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# REVIEW OF PARTICLE PHYSICS\*

Particle Data Group

## *Abstract*

The *Review* summarizes much of particle physics and cosmology. Using data from previous editions, plus 3,062 new measurements from 721 papers, we list, evaluate, and average measured properties of gauge bosons and the recently discovered Higgs boson, leptons, quarks, mesons, and baryons. We summarize searches for hypothetical particles such as supersymmetric particles, heavy bosons, axions, dark photons, etc. All the particle properties and search limits are listed in Summary Tables. We also give numerous tables, figures, formulae, and reviews of topics such as Higgs Boson Physics, Supersymmetry, Grand Unified Theories, Neutrino Mixing, Dark Energy, Dark Matter, Cosmology, Particle Detectors, Colliders, Probability and Statistics. Among the 117 reviews are many that are new or heavily revised, including new reviews on Pentaquarks and Inflation.

The complete *Review* is published online in a journal and on the website of the Particle Data Group (<http://pdg.lbl.gov>). The printed *PDG Book* contains the Summary Tables and all review articles but no longer includes the detailed tables from the Particle Listings. A *Booklet* with the Summary Tables and abbreviated versions of some of the review articles is also available.

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## HIGHLIGHTS OF THE 2016 EDITION OF THE REVIEW OF PARTICLE PHYSICS

721 new papers with 3062 new measurements

- Over 332 new papers from **LHC** experiments (ATLAS, CMS, and LHCb).
- Extensive up-to-date **Higgs boson** coverage from 79 new papers with 172 measurements.
- **Supersymmetry**: 82 new papers with major exclusions.
- **Top quark**: 55 new papers.
- Latest from ***B*-meson** physics: 133 papers with 542 measurements.
- New  $\tau$  **branching fractions** fit in collaboration with the HFAG-Tau group.
- New limits on neutrinoless **double- $\beta$  decays**.
- Updated and new results in **neutrino mixing** on  $\Delta m^2$  and mixing angle measurements.
- Experimental Tests of Gravitational Theory review includes LIGO observation of **gravitational waves**.
- Cosmology reviews updated to include **2015 Planck** results.
- **Periodic Table** 7th row completed; significantly revised Atomic-Nuclear Properties website.

117 reviews (most are revised)

- **New reviews on:**
  - Inflation
  - Pentaquarks
  - Pole Structure of the  $\Lambda(1405)$  Region
- **Significant update/revision** to reviews on:
  - Higgs Boson Physics
  - Grand Unified Theories
  - Dark Energy, Dark Matter and CMB
  - Cosmological Parameters, Astrophysical Constants and Parameters
  - Neutrino Mass, Mixing, and Flavor Change
  - Neutrino Cross Section Measurements
  - $W'$  and  $Z'$  bosons searches
  - Searches for Quark and Lepton Compositeness
  - Leptonic Decays of Charged Pseudoscalar Mesons
  - Particle Detectors for accelerator and non-accelerator physics, including new section on Accelerator Neutrino Detectors
  - High-Energy Collider Parameters

See [pdgLive.lbl.gov](http://pdgLive.lbl.gov) for online access to PDG database.

See [pdg.lbl.gov/AtomicNuclearProperties](http://pdg.lbl.gov/AtomicNuclearProperties) for Atomic Properties of Materials.

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\*The divider sheets give more detailed indices for each main section of the Particle Listings.



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## INTRODUCTION

### 1. Overview

The *Review of Particle Physics* is a review of the field of Particle Physics and of related areas in Cosmology. It consists of “Summary Tables”, “Particle Listings”, and “Reviews, Tables, and Plots”. The latter covers a wide variety of theoretical and experimental topics and provides a quick reference for the practicing particle physicist.

The Summary Tables give our best values and limits for particle properties such as masses, widths or lifetimes, and branching fractions, as well as an extensive summary of searches for hypothetical particles and a summary of experimental tests of conservation laws.

The Particle Listings are a compilation/evaluation of data on particle properties. They contain all the data used to get the values given in the Summary Tables. The Particle Listings also give information on unconfirmed particles and on particle searches, as well as reviews on subjects of particular interest or controversy. In this edition, the Particle Listings include 3,062 new measurements from 721 papers, in addition to the 35,436 measurements from 9,843 papers that first appeared in previous editions [1]. Because of the large quantity of data, the Particle Listings are not an archive of all published data on particle properties. We refer interested readers to earlier editions for data now considered to be obsolete.

The book version of the *Review* is published in even-numbered years. This edition is an updating through January 2016 (and, in some areas, well into 2016). The content of this *Review* is available on the web and is updated between printed editions.

We organize the particles into six categories:

Gauge and Higgs bosons  
Leptons  
Quarks  
Mesons  
Baryons  
Searches for monopoles, supersymmetry,  
compositeness, extra dimensions, *etc.*

The last category only includes searches for particles that do not belong to the previous groups; searches for heavy charged leptons and massive neutrinos, by contrast, are with the leptons.

In Sec. 2 of this Introduction, we list the main areas of responsibility of the authors of the Particle Listings. Our many consultants, without whom we would not have been able to produce this *Review*, are acknowledged in Sec. 3. In Sec. 4, we mention briefly the naming scheme for hadrons. In Sec. 5, we discuss our procedures for choosing among measurements of particle properties and for obtaining best values of the properties from the measurements.

The accuracy and usefulness of this *Review* depend in large part on interaction between its users and the authors. We appreciate comments, criticisms, and suggestions for improvements of any kind. Please send them to the appropriate author, according to the list of responsibilities in Sec. 2 below, or to [pdg@lbl.gov](mailto:pdg@lbl.gov).

The complete *Review* is published online in a journal and on the PDG website (<http://pdg.lbl.gov>). In addition to the online publication, the Review is available in different formats:

- The printed *PDG Book* contains the Summary Tables and all review articles. In contrast to previous editions, the detailed tables from the Particle Listings are no longer printed.
- The *Particle Physics Booklet* includes the Summary Tables and abbreviated versions of some of the review articles in a pocket format.
- *pdgLive* (<http://pdgLive.lbl.gov>) is a web application for online access to the PDG database.
- Files that can be downloaded from the PDG website include a table of masses, widths, and PDG Monte Carlo particle ID numbers; PDF files of the entire PDG Book and Booklet; individual review articles; all figures; and an archive file containing the complete PDG website (except for *pdgLive*).

Copies of the *PDG Book* or the *Particle Physics Booklet* can be ordered from our website or directly at <http://pdg.lbl.gov/order>. For special requests only, please email [pdg@lbl.gov](mailto:pdg@lbl.gov) in North and South America, Australia, and the Far East, and [pdg-products@cern.ch](mailto:pdg-products@cern.ch) in all other areas.

### 2. Particle Listings responsibilities

\* Asterisk indicates the people to contact with questions or comments about Particle Listings sections.

#### Gauge and Higgs bosons

$\gamma$	C. Grab, D.E. Groom*
Gluons	R.M. Barnett,* A.V. Manohar
Graviton	D.E. Groom*
$W, Z$	A. Gurtu,* M. Grünewald*
Higgs bosons	K. Hikasa, G. Weiglein*
Heavy bosons	S. Pagan Griso,* M. Tanabashi
Axions	K.A. Olive, F. Takahashi, G. Raffelt*

#### Leptons

Neutrinos	M. Goodman, C.-J. Lin,* K. Nakamura, K.A. Olive, A. Piepke, P. Vogel
$e, \mu$	C. Grab, C.-J. Lin*
$\tau$	K.G. Hayes, K. Mönig*

#### Quarks

Quarks	R.M. Barnett,* A.V. Manohar
Top quark	R.M. Barnett,* Y. Sumino
$b', t'$	R.M. Barnett,* Y. Sumino
Free quark	S. Pagan Griso*

#### Mesons

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Unstable mesons	C. Amsler, M. Doser,* S. Eidelman,* T. Gutsche, C. Hanhart, B. Heltsley, J.J. Hernández-Rey, A. Masoni, R.E. Mitchell, S. Navas, C. Patrignani, S. Spanier, N.A. Törnqvist, G. Venanzoni
$K$ (stable)	G. D'Ambrosio, C.-J. Lin*
$D$ (stable, no mix.)	J. Rademacker, C.G. Wohl*
$D^0$ mixing	D.M. Asner, W.-M. Yao*

Baryons

$B$ (stable)	A. Cerri,* P. Eerola, M. Kreps, Y. Kwon, W.-M. Yao*
Stable baryons	C. Grab, C.G. Wohl*
Unstable baryons	V. Burkert, E. Klempt, M. Pennington, L. Tiator, R.L. Workman*
Charmed baryons	J. Rademacker, C.G. Wohl*
Bottom baryons	A. Cerri,* P. Eerola, M. Kreps, Y. Kwon, W.-M. Yao*

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Technicolor	K. Agashe,* M. Tanabashi
Compositeness	M. Tanabashi, J. Terning*
Extra Dimensions	D.A. Dwyer,* T. Gherghetta
WIMPs and Other	K. Hikasa*

**3. Consultants**

The Particle Data Group benefits greatly from the assistance of some 700 physicists who are asked to verify every piece of data entered into this *Review*. Of special value is the advice of the PDG Advisory Committee which meets biennially and thoroughly reviews all aspects of our operation. The members of the 2016 committee are:

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J. Olson (Princeton)  
A. Slosar (BNL)  
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#### 4. Naming scheme for hadrons

We introduced in the 1986 edition [2] a new naming scheme for the hadrons. Changes from older terminology affected mainly the heavier mesons made of  $u$ ,  $d$ , and  $s$  quarks. Otherwise, the only important change to known hadrons was that the  $F^\pm$  became the  $D_s^\pm$ . None of the lightest pseudoscalar or vector mesons changed names, nor did the  $c\bar{c}$  or  $b\bar{b}$  mesons (we do, however, now use  $\chi_c$  for the  $c\bar{c}$   $\chi$  states), nor did any of the established baryons. The Summary Tables give both the new and old names whenever a change has occurred.

The scheme is described in “Naming Scheme for Hadrons” (p. 130) of this *Review*.

We give here our conventions on type-setting style. Particle symbols are italic (or slanted) characters:  $e^-$ ,  $p$ ,  $\Lambda$ ,  $\pi^0$ ,  $K_L$ ,  $D_s^+$ ,  $b$ . Charge is indicated by a superscript:  $B^-$ ,  $\Delta^{++}$ . Charge is not normally indicated for  $p$ ,  $n$ , or the quarks, and is optional for neutral isosinglets:  $\eta$  or  $\eta^0$ . Antiparticles and particles are distinguished by charge for charged leptons and mesons:  $\tau^+$ ,  $K^-$ . Otherwise, distinct antiparticles are indicated by a bar (overline):  $\bar{\nu}_\mu$ ,  $\bar{t}$ ,  $\bar{p}$ ,  $\bar{K}^0$ , and  $\bar{\Sigma}^+$  (the antiparticle of the  $\Sigma^-$ ).

#### 5. Procedures

**5.1. Selection and treatment of data :** The Particle Listings contain all relevant data known to us that are published in journals. With very few exceptions, we do not include results from preprints or conference reports. Nor do we include data that are of historical importance only (the Listings are not an archival record). We search every volume of 20 journals through our cutoff date for relevant data. We also include later published papers that are sent to us by the authors (or others).

In the Particle Listings, we clearly separate measurements that are used to calculate or estimate values given in the Summary Tables from measurements that are not used. We give explanatory comments in many such cases. Among the reasons a measurement might be excluded are the following:

- It is superseded by or included in later results.
- No error is given.
- It involves assumptions we question.
- It has a poor signal-to-noise ratio, low statistical significance, or is otherwise of poorer quality than other data available.
- It is clearly inconsistent with other results that appear to be more reliable. Usually we then state the criterion, which sometimes is quite subjective, for selecting “more reliable” data for averaging. See Sec. 5.4.
- It is not independent of other results.
- It is not the best limit (see below).
- It is quoted from a preprint or a conference report.

In some cases, *none* of the measurements is entirely reliable and no average is calculated. For example, the masses of many of the baryon resonances, obtained from partial-wave analyses, are quoted as estimated ranges thought to probably include the true values, rather than as averages with errors. This is discussed in the Baryon Particle Listings.

For upper limits, we normally quote in the Summary Tables the strongest limit. We do not average or combine upper limits except in a very few cases where they may be re-expressed as measured numbers with Gaussian errors.

As is customary, we assume that particle and antiparticle share the same spin, mass, and mean life. The Tests of Conservation Laws table, following the Summary Tables, lists tests of *CPT* as well as other conservation laws.

We use the following indicators in the Particle Listings to tell how we get values from the tabulated measurements:

- OUR AVERAGE—From a weighted average of selected data.
- OUR FIT—From a constrained or overdetermined multi-parameter fit of selected data.
- OUR EVALUATION—Not from a direct measurement, but evaluated from measurements of related quantities.
- OUR ESTIMATE—Based on the observed range of the data. Not from a formal statistical procedure.
- OUR LIMIT—For special cases where the limit is evaluated by us from measured ratios or other data. Not from a direct measurement.

An experimentalist who sees indications of a particle will of course want to know what has been seen in that region in the past. Hence we include in the Particle Listings all reported states that, in our opinion, have sufficient statistical merit and that have not been disproved by more reliable data. However, we promote to the Summary Tables only those states that we feel are well established. This judgment is, of course, somewhat subjective and no precise criteria can be given. For more detailed discussions, see the minireviews in the Particle Listings.

**5.2. Averages and fits:** We divide this discussion on obtaining averages and errors into three sections: (1) treatment of errors; (2) unconstrained averaging; (3) constrained fits.

**5.2.1. Treatment of errors:** In what follows, the “error”  $\delta x$  means that the range  $x \pm \delta x$  is intended to be a 68.3% confidence interval about the central value  $x$ . We treat this error as if it were Gaussian. Thus when the error is Gaussian,  $\delta x$  is the usual one standard deviation ( $1\sigma$ ). Many experimenters now give statistical and systematic errors separately, in which case we usually quote both errors, with the statistical error first. For averages and fits, we then add the two errors in quadrature and use this combined error for  $\delta x$ .

When experimenters quote asymmetric errors  $(\delta x)^+$  and  $(\delta x)^-$  for a measurement  $x$ , the error that we use for that measurement in making an average or a fit with other measurements is a continuous function of these three quantities. When the resultant average or fit  $\bar{x}$  is less than  $x - (\delta x)^-$ , we use  $(\delta x)^-$ ; when it is greater than  $x + (\delta x)^+$ , we use  $(\delta x)^+$ . In between, the error we use is a linear function of  $x$ . Since the errors we use are functions of the result, we iterate to get the final result. Asymmetric output errors are

determined from the input errors assuming a linear relation between the input and output quantities.

In fitting or averaging, we usually do not include correlations between different measurements, but we try to select data in such a way as to reduce correlations. Correlated errors are, however, treated explicitly when there are a number of results of the form  $A_i \pm \sigma_i \pm \Delta$  that have identical systematic errors  $\Delta$ . In this case, one can first average the  $A_i \pm \sigma_i$  and then combine the resulting statistical error with  $\Delta$ . One obtains, however, the same result by averaging  $A_i \pm (\sigma_i^2 + \Delta_i^2)^{1/2}$ , where  $\Delta_i = \sigma_i \Delta [\sum (1/\sigma_j^2)]^{1/2}$ . This procedure has the advantage that, with the modified systematic errors  $\Delta_i$ , each measurement may be treated as independent and averaged in the usual way with other data. Therefore, when appropriate, we adopt this procedure. We tabulate  $\Delta$  and invoke an automated procedure that computes  $\Delta_i$  before averaging and we include a note saying that there are common systematic errors.

Another common case of correlated errors occurs when experimenters measure two quantities and then quote the two and their difference, *e.g.*,  $m_1$ ,  $m_2$ , and  $\Delta = m_2 - m_1$ . We cannot enter all of  $m_1$ ,  $m_2$  and  $\Delta$  into a constrained fit because they are not independent. In some cases, it is a good approximation to ignore the quantity with the largest error and put the other two into the fit. However, in some cases correlations are such that the errors on  $m_1$ ,  $m_2$  and  $\Delta$  are comparable and none of the three values can be ignored. In this case, we put all three values into the fit and invoke an automated procedure to increase the errors prior to fitting such that the three quantities can be treated as independent measurements in the constrained fit. We include a note saying that this has been done.

**5.2.2. Unconstrained averaging:** To average data, we use a standard weighted least-squares procedure and in some cases, discussed below, increase the errors with a “scale factor.” We begin by assuming that measurements of a given quantity are uncorrelated, and calculate a weighted average and error as

$$\bar{x} \pm \delta\bar{x} = \frac{\sum_i w_i x_i}{\sum_i w_i} \pm (\sum_i w_i)^{-1/2}, \quad (1)$$

where

$$w_i = 1/(\delta x_i)^2.$$

Here  $x_i$  and  $\delta x_i$  are the value and error reported by the  $i$ th experiment, and the sums run over the  $N$  experiments. We then calculate  $\chi^2 = \sum w_i (\bar{x} - x_i)^2$  and compare it with  $N - 1$ , which is the expectation value of  $\chi^2$  if the measurements are from a Gaussian distribution.

If  $\chi^2/(N - 1)$  is less than or equal to 1, and there are no known problems with the data, we accept the results.

If  $\chi^2/(N - 1)$  is very large, we may choose not to use the average at all. Alternatively, we may quote the calculated average, but then make an educated guess of the error, a conservative estimate designed to take into account known problems with the data.

Finally, if  $\chi^2/(N - 1)$  is greater than 1, but not greatly so, we still average the data, but then also do the following:

(a) We increase our quoted error,  $\delta\bar{x}$  in Eq. (1), by a scale factor  $S$  defined as

$$S = [\chi^2/(N - 1)]^{1/2}. \quad (2)$$

Our reasoning is as follows. The large value of the  $\chi^2$  is likely to be due to underestimation of errors in at least one of the experiments. Not knowing which of the errors are underestimated, we assume they are all underestimated by the same factor  $S$ . If we scale up all the input errors by this factor, the  $\chi^2$  becomes  $N - 1$ , and of course the output error  $\delta\bar{x}$  scales up by the same factor. See Ref. 3.

When combining data with widely varying errors, we modify this procedure slightly. We evaluate  $S$  using only the experiments with smaller errors. Our cutoff or ceiling on  $\delta x_i$  is arbitrarily chosen to be

$$\delta_0 = 3N^{1/2} \delta\bar{x},$$

where  $\delta\bar{x}$  is the unscaled error of the mean of all the experiments. Our reasoning is that although the low-precision experiments have little influence on the values  $\bar{x}$  and  $\delta\bar{x}$ , they can make significant contributions to the  $\chi^2$ , and the contribution of the high-precision experiments thus tends to be obscured. Note that if each experiment has the same error  $\delta x_i$ , then  $\delta\bar{x}$  is  $\delta x_i/N^{1/2}$ , so each  $\delta x_i$  is well below the cutoff. (More often, however, we simply exclude measurements with relatively large errors from averages and fits: new, precise data chase out old, imprecise data.)

Our scaling procedure has the property that if there are two values with comparable errors separated by much more than their stated errors (with or without a number of other values of lower accuracy), the scaled-up error  $\delta\bar{x}$  is approximately half the interval between the two discrepant values.

We emphasize that our scaling procedure for *errors* in no way affects central values. And if you wish to recover the unscaled error  $\delta\bar{x}$ , simply divide the quoted error by  $S$ .

(b) If the number  $M$  of experiments with an error smaller than  $\delta_0$  is at least three, and if  $\chi^2/(M - 1)$  is greater than 1.25, we show in the Particle Listings an ideogram of the data. Figure 1 is an example. Sometimes one or two data points lie apart from the main body; other times the data split into two or more groups. We extract no numbers from these ideograms; they are simply visual aids, which the reader may use as he or she sees fit.

Each measurement in an ideogram is represented by a Gaussian with a central value  $x_i$ , error  $\delta x_i$ , and area proportional to  $1/\delta x_i$ . The choice of  $1/\delta x_i$  for the area is somewhat arbitrary. With this choice, the center of gravity of the ideogram corresponds to an average that uses weights  $1/\delta x_i$  rather than the  $(1/\delta x_i)^2$  actually used in the averages. This may be appropriate when some of the experiments have seriously underestimated systematic errors. However, since for this choice of area the height of the Gaussian for each measurement is proportional to  $(1/\delta x_i)^2$ , the peak position of the ideogram will often favor the high-precision measurements at least as much as does the least-squares average. See our 1986 edition [2] for a detailed discussion of the use of ideograms.

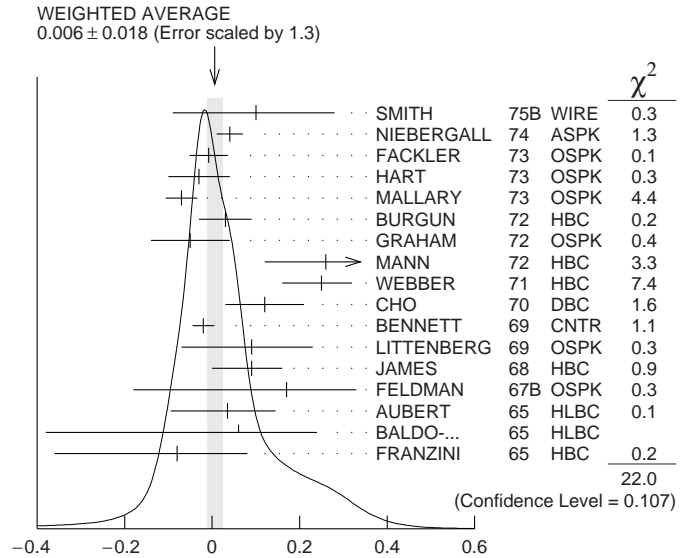
**5.2.3. Constrained fits:** In some cases, such as branching ratios or masses and mass differences, a constrained fit may be needed to obtain the best values of a set of parameters. For example, most branching ratios and rate measurements are analyzed by making a simultaneous least-squares fit to all the data and extracting the partial decay fractions  $P_i$ , the partial widths  $\Gamma_i$ , the full width  $\Gamma$  (or mean life), and the associated error matrix.

Assume, for example, that a state has  $m$  partial decay fractions  $P_i$ , where  $\sum P_i = 1$ . These have been measured in  $N_r$  different ratios  $R_r$ , where, e.g.,  $R_1 = P_1/P_2$ ,  $R_2 = P_1/P_3$ , etc. [We can handle any ratio  $R$  of the form  $\sum \alpha_i P_i / \sum \beta_i P_i$ , where  $\alpha_i$  and  $\beta_i$  are constants, usually 1 or 0. The forms  $R = P_i P_j$  and  $R = (P_i P_j)^{1/2}$  are also allowed.] Further assume that *each* ratio  $R$  has been measured by  $N_k$  experiments (we designate each experiment with a subscript  $k$ , e.g.,  $R_{1k}$ ). We then find the best values of the fractions  $P_i$  by minimizing the  $\chi^2$  as a function of the  $m - 1$  independent parameters:

$$\chi^2 = \sum_{r=1}^{N_r} \sum_{k=1}^{N_k} \left( \frac{R_{rk} - R_r}{\delta R_{rk}} \right)^2, \quad (3)$$

where the  $R_{rk}$  are the measured values and  $R_r$  are the fitted values of the branching ratios.

In addition to the fitted values  $\bar{P}_i$ , we calculate an error matrix  $\langle \delta \bar{P}_i \delta \bar{P}_j \rangle$ . We tabulate the diagonal elements of  $\delta \bar{P}_i = \langle \delta \bar{P}_i \delta \bar{P}_i \rangle^{1/2}$  (except that some errors are scaled as discussed below). In the Particle Listings, we give the complete correlation matrix; we also calculate the fitted value of each ratio, for comparison with the input data, and list it above the relevant input, along with a simple unconstrained average of the same input.



**Figure 1:** A typical ideogram. The arrow at the top shows the position of the weighted average, while the width of the shaded pattern shows the error in the average after scaling by the factor  $S$ . The column on the right gives the  $\chi^2$  contribution of each of the experiments. Note that the next-to-last experiment, denoted by the incomplete error flag ( $\perp$ ), is not used in the calculation of  $S$  (see the text).



Three comments on the example above:

(1) There was no connection assumed between measurements of the full width and the branching ratios. But often we also have information on partial widths  $\Gamma_i$  as well as the total width  $\Gamma$ . In this case we must introduce  $\Gamma$  as a parameter in the fit, along with the  $P_i$ , and we give correlation matrices for the widths in the Particle Listings.

(2) We try to pick those ratios and widths that are as independent and as close to the original data as possible. When one experiment measures all the branching fractions and constrains their sum to be one, we leave one of them (usually the least well-determined one) out of the fit to make the set of input data more nearly independent. We now do allow for correlations between input data.

(3) We calculate scale factors for both the  $R_r$  and  $P_i$  when the measurements for any  $R$  give a larger-than-expected contribution to the  $\chi^2$ . According to Eq. (3), the double sum for  $\chi^2$  is first summed over experiments  $k = 1$  to  $N_k$ , leaving a single sum over ratios  $\chi^2 = \sum \chi_r^2$ . One is tempted to define a scale factor for the ratio  $r$  as  $S_r^2 = \chi_r^2 / \langle \chi_r^2 \rangle$ . However, since  $\langle \chi_r^2 \rangle$  is not a fixed quantity (it is somewhere between  $N_k$  and  $N_{k-1}$ ), we do not know how to evaluate this expression. Instead we define

$$S_r^2 = \frac{1}{N_k} \sum_{k=1}^{N_k} \frac{(R_{rk} - \bar{R}_r)^2}{\langle (R_{rk} - \bar{R}_r)^2 \rangle}. \quad (4)$$

With this definition the expected value of  $S_r^2$  is one. We can show that

$$\langle (R_{rk} - \bar{R}_r)^2 \rangle = \langle (\delta R_{rk})^2 \rangle - (\delta \bar{R}_r)^2, \quad (5)$$

where  $\delta \bar{R}_r$  is the fitted error for ratio  $r$ .

The fit is redone using errors for the branching ratios that are scaled by the larger of  $S_r$  and unity, from which new and often larger errors  $\delta \bar{P}_i'$  are obtained. The scale factors we finally list in such cases are defined by  $S_i = \delta \bar{P}_i' / \delta \bar{P}_i$ . However, in line with our policy of not letting  $S$  affect the central values, we give the values of  $\bar{P}_i$  obtained from the original (unscaled) fit.

There is one special case in which the errors that are obtained by the preceding procedure may be changed. When a fitted branching ratio (or rate)  $\bar{P}_i$  turns out to be less than three standard deviations ( $\delta \bar{P}_i'$ ) from zero, a new smaller error ( $\delta \bar{P}_i''$ ) is calculated on the low side by requiring the area under the Gaussian between  $\bar{P}_i - (\delta \bar{P}_i'')$  and  $\bar{P}_i$  to be 68.3% of the area between zero and  $\bar{P}_i$ . A similar correction is made for branching fractions that are within three standard deviations of one. This keeps the quoted errors from overlapping the boundary of the physical region.

**5.3. Rounding:** While the results shown in the Particle Listings are usually exactly those published by the experiments, the numbers that appear in the Summary Tables (means, averages and limits) are subject to a set of rounding rules.

The basic rule states that if the three highest order digits of the error lie between 100 and 354, we round to two significant digits. If they lie between 355 and 949, we round to one significant digit. Finally, if they lie between 950 and 999, we round up to 1000 and keep two significant digits. In all cases, the central value is given with a precision

that matches that of the error. So, for example, the result (coming from an average)  $0.827 \pm 0.119$  would appear as  $0.83 \pm 0.12$ , while  $0.827 \pm 0.367$  would turn into  $0.8 \pm 0.4$ .

Rounding is not performed if a result in a Summary Table comes from a single measurement, without any averaging. In that case, the number of digits published in the original paper is kept, unless we feel it inappropriate. Note that, even for a single measurement, when we combine statistical and systematic errors in quadrature, rounding rules apply to the result of the combination. It should be noted also that most of the limits in the Summary Tables come from a single source (the best limit) and, therefore, are not subject to rounding.

Finally, we should point out that in several instances, when a group of results come from a single fit to a set of data, we have chosen to keep two significant digits for all the results. This happens, for instance, for several properties of the  $W$  and  $Z$  bosons and the  $\tau$  lepton.

**5.4. Discussion:** The problem of averaging data containing discrepant values is nicely discussed by Taylor in Ref. 4. He considers a number of algorithms that attempt to incorporate inconsistent data into a meaningful average. However, it is difficult to develop a procedure that handles simultaneously in a reasonable way two basic types of situations: (a) data that lie apart from the main body of the data are incorrect (contain unreported errors); and (b) the opposite—it is the main body of data that is incorrect. Unfortunately, as Taylor shows, case (b) is not infrequent. He concludes that the choice of procedure is less significant than the initial choice of data to include or exclude.

We place much emphasis on this choice of data. Often we solicit the help of outside experts (consultants). Sometimes, however, it is simply impossible to determine which of a set of discrepant measurements are correct. Our scale-factor technique is an attempt to address this ignorance by increasing the error. In effect, we are saying that present experiments do not allow a precise determination of this quantity because of unresolvable discrepancies, and one must await further measurements. The reader is warned of this situation by the size of the scale factor, and if he or she desires can go back to the literature (via the Particle Listings) and redo the average with a different choice of data.

Our situation is less severe than most of the cases Taylor considers, such as estimates of the fundamental constants like  $\hbar$ , etc. Most of the errors in his case are dominated by systematic effects. For our data, statistical errors are often at least as large as systematic errors, and statistical errors are usually easier to estimate. A notable exception occurs in partial-wave analyses, where different techniques applied to the same data yield different results. In this case, as stated earlier, we often do not make an average but just quote a range of values.

A brief history of early Particle Data Group averages is given in Ref. 3. Figure 2 shows some histories of our values of a few particle properties. Sometimes large changes occur. These usually reflect the introduction of significant new data or the discarding of older data. Older data are discarded in favor of newer data when it is felt that the newer data have smaller systematic errors, or have more checks on systematic errors, or have made corrections unknown at the time of the older experiments, or simply have much smaller errors. Sometimes, the scale factor becomes large near the time at which a large jump takes place, reflecting

the uncertainty introduced by the new and inconsistent data. By and large, however, a full scan of our history plots shows a dull progression toward greater precision at central values quite consistent with the first data points shown.

We conclude that the reliability of the combination of experimental data and our averaging procedures is usually good, but it is important to be aware that fluctuations outside of the quoted errors can and do occur.

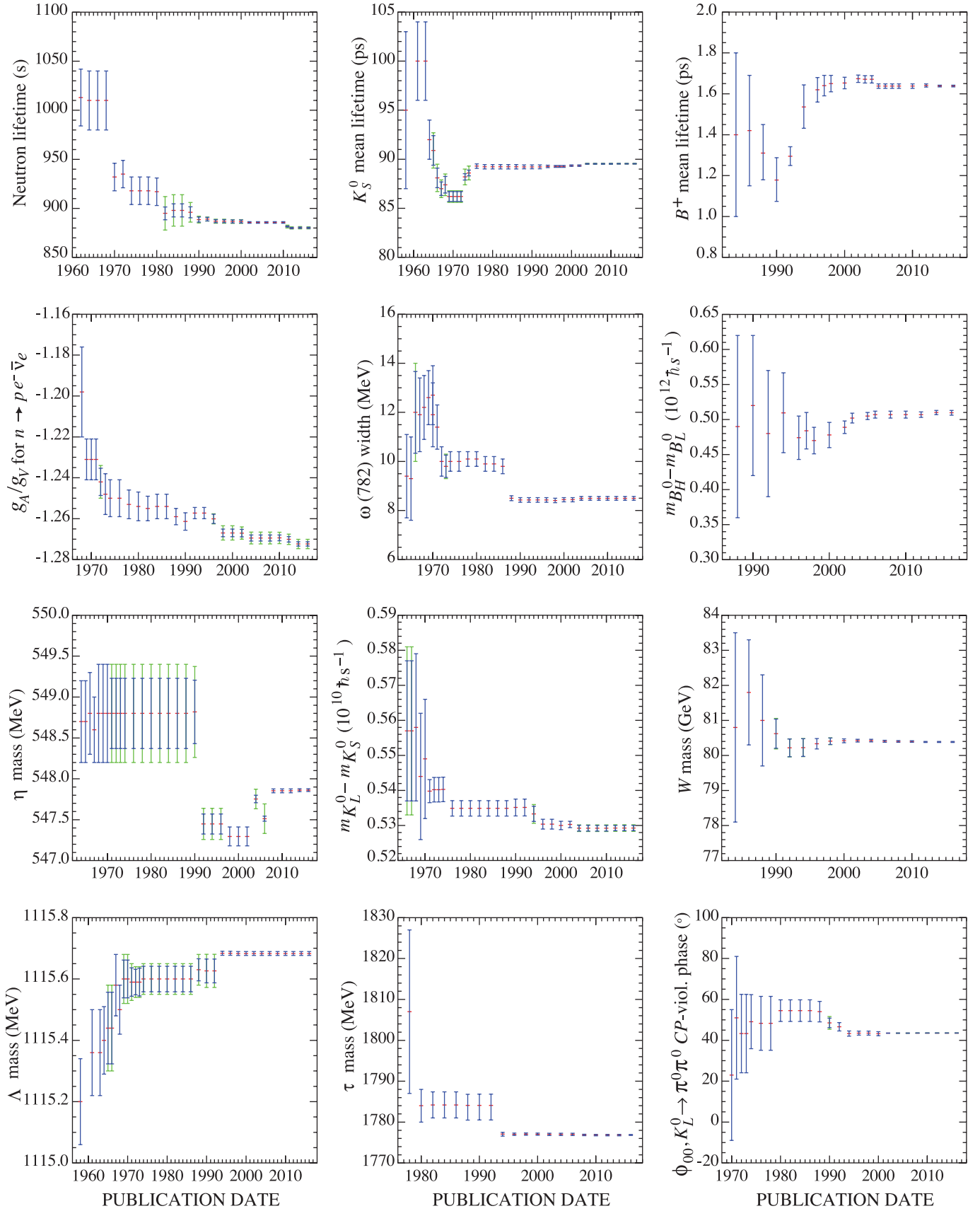
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## REFERENCES

1. The previous edition was Particle Data Group: K.A. Olive *et al.*, Chin. Phys. C **38**, 090001 (2014).
2. Particle Data Group: M. Aguilar-Benitez *et al.*, Phys. Lett. **170B** (1986).
3. A.H. Rosenfeld, Ann. Rev. Nucl. Sci. **25**, 555 (1975).
4. B.N. Taylor, “Numerical Comparisons of Several Algorithms for Treating Inconsistent Data in a Least-Squares Adjustment of the Fundamental Constants,” U.S. National Bureau of Standards NBSIR 81-2426 (1982).



**Figure 2:** A historical perspective of values of a few particle properties tabulated in this *Review* as a function of date of publication of the *Review*. A full error bar indicates the quoted error; a thick-lined portion indicates the same but without the “scale factor.”

# ONLINE PARTICLE PHYSICS INFORMATION

Updated Nov. 2015 by T. Basaglia (CERN), A. Holtkamp (CERN).<sup>†</sup>

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## 1. Introduction

The collection of online information resources in particle physics and related areas presented in this chapter is of necessity incomplete. An expanded and regularly updated online version can be found at:

[http://library.web.cern.ch/particle\\_physics\\_information](http://library.web.cern.ch/particle_physics_information)

Suggestions for additions and updates are very welcome.<sup>†</sup>

## 2. Particle Data Group (PDG) resources

- **Review of Particle Physics (RPP)** A comprehensive report on the fields of particle physics and related areas of cosmology and astrophysics, including both review articles and a compilation/evaluation of data on particle properties. The review section includes articles, tables and plots on a wide variety of theoretical and experimental topics of interest to particle physicists and astrophysicists. The particle properties section provides tables of published measurements as well as the Particle Data Groups best values and limits for particle properties such as masses, widths, lifetimes, and branching fractions, and an extensive summary of searches for hypothetical particles. RPP is published as a 1500-page book every two years, with partial updates made available once each year on the web.

All the contents of the book version of RPP are available online:

<http://pdg.lbl.gov>

The printed book can be ordered:

[http://pdg.lbl.gov/2015/html/receive\\_our\\_products.html](http://pdg.lbl.gov/2015/html/receive_our_products.html)

Of historical interest is the complete RPP collection which can be found online:

[http://library.web.cern.ch/PDG\\_publications/review\\_particle\\_physics](http://library.web.cern.ch/PDG_publications/review_particle_physics)

- **Particle Physics booklet:** An abridged version of the Review of Particle Physics available as a pocket-sized 300-page booklet. Although produced in print and available online only as a PDF file, the booklet is included in this guide because it is one of the most useful summaries of physics data. The booklet contains an

abbreviated set of reviews and the summary tables from the most recent edition of the Review of Particle Physics.

The PDF file of the booklet can be downloaded:

<http://pdg.lbl.gov/current/booklet.pdf>

The printed booklet can be ordered:

[http://pdg.lbl.gov/2015/html/receive\\_our\\_products.html](http://pdg.lbl.gov/2015/html/receive_our_products.html)

- **PDGLive:** A web application for browsing the contents of the PDG database that contains the information published in the Review of Particle Physics. It allows one to navigate to a particle of interest, see a summary of the information available, and then proceed to the detailed information published in the Review of Particle Physics. Data entries are directly linked to the corresponding bibliographic information in INSPIRE.

<http://pdglive.lbl.gov>

- **Computer-readable files:** Data files that can be downloaded from PDG include tables of particle masses and widths, PDG Monte Carlo particle numbers, and cross-section data. The files are updated with each new edition of the Review of Particle Physics.

[http://pdg.lbl.gov/current/html/computer\\_read.html](http://pdg.lbl.gov/current/html/computer_read.html)

## 3. Particle Physics Information Platforms

- **INSPIRE:** The time-honored SPIRES database suite has in November 2011 been replaced by INSPIRE, which combines the most successful aspects of SPIRES - like comprehensive content and high-quality metadata - with the modern technology of Invenio, the CERN open-source digital-library software, offering major improvements like increased speed and Google-like free-text search syntax. INSPIRE serves as one-stop information platform for the particle physics community, comprising 8 interlinked databases on literature, conferences, institutions, journals, researchers, experiments, jobs and data. INSPIRE is jointly developed and maintained by CERN, DESY, Fermilab, IHEP and SLAC. Close interaction with the user community and with arXiv, ADS, HepData, PDG and publishers is the backbone of INSPIRE's evolution.

<http://inspirehep.net/>

INSPIRE is integrated with ORCID (Open Researcher and Contributor ID), a persistent identifier that enables researchers to connect services and get credit for their works.

<http://orcid.org/>

INSPIRE is currently developing a new version of the portal, maintaining its quality standards and introducing new functionality. The INSPIRE Labs site is available at:

<http://labs.inspirehep.net>

blog: <http://blog.inspirehep.net/>

twitter: @inspirehep

## 4. Literature Databases

- **ADS:** The SAO/NASA Astrophysics Data System is a Digital Library portal offering access to 11 million bibliographic records in Astronomy and Physics. The ADS's search engine also indexes the full-text for approximately four million publications in this collection and tracks citations, which now amount to over 80 million links. The system also provides access and links to a wealth of external resources, including electronic articles hosted by publishers and arXiv, data catalogs and a variety of data products hosted by the astronomy archives worldwide. The ADS can be accessed at

<http://ads.harvard.edu/>

- **arXiv.org:** A repository of full text papers in physics, mathematics, computer science, statistics, nonlinear sciences, quantitative finance and quantitative biology interlinked with ADS and INSPIRE. Papers are usually submitted by their authors to arXiv in advance of submission to a journal for publication. Primarily covers 1991 to the present but authors are encouraged to post older papers retroactively. Permits searching by author, title, and words in abstract and experimentally also in the fulltext. Allows limiting by subfield archive or by date. Daily update alerts by subfield are available by email and RSS.

<sup>†</sup> Please send comments and corrections to  
Annette.Holtkamp@cern.ch.

<http://arXiv.org>

- **CDS:** The CERN Document Server contains records of more than 1,000,000 CERN and non-CERN articles, preprints, theses. It includes records for internal and technical notes, official CERN committee documents, and multimedia objects. CDS is going to focus on its role as institutional repository covering all CERN material from the early 50s and reflecting the holdings of the CERN library. Non-CERN particle and accelerator physics content is in the process of being exported to INSPIRE.

<http://cds.cern.ch>

- **INSPIRE HEP:** The HEP collection, the flagship of the INSPIRE suite, serves more than 1.1 million bibliographic records with a growing number of fulltexts attached and metadata including author affiliations, abstracts, references, experiments, keywords as well as links to arXiv, ADS, PDG, HepData and publisher platforms. It provides fast metadata and fulltext searches, plots extracted from fulltext, author disambiguation, author profile pages and citation analysis and is expanding its content to, e.g., experimental notes.

<http://inspirehep.net>

- **JACoW:** The Joint Accelerator Conference Website publishes the proceedings of APAC, EPAC, PAC, IPAC, ABDW, BIW, COOL, CYCLOTRONS, DIPAC, ECRIS, FEL, HIAT, ICALEPCS, IBIC, ICAP, LINAC, North American PAC, PCaPAC, RuPAC, SRF. A custom interface allows searching on keywords, titles, authors, and in the fulltext.

<http://www.jacow.org/>

- **KISS (KEK Information Service System) for preprints:** The KEK Library preprint and technical report database contains bibliographic records of preprints and technical reports held in the KEK library with links to the full text images of more than 100,000 papers scanned from their worldwide collection of preprints. Particularly useful for older scanned preprints. KISS links are included in INSPIRE HEP.

[http://www-lib.kek.jp/KISS/kiss\\_preprint.html](http://www-lib.kek.jp/KISS/kiss_preprint.html)

- **MathSciNet:** This database of almost 3 million items provides reviews, abstracts and bibliographic information for much of the mathematical sciences literature. Over 100,000 new items are added each year, most of them classified according to the Mathematics Subject Classification. Authors are uniquely identified, enabling a search for publications by individual author. Over 80,000 reviews on the current published literature are added each year. Citation data allows to track the history and influence of research publications.

<http://www.ams.org/mathscinet>

- **OSTI SciTech Connect:** A portal to free, publicly available DOE-sponsored R&D results including technical reports, bibliographic citations, journal articles, conference papers, books, multimedia and data information. SciTech Connect is a consolidation of two core DOE search engines, the Information Bridge and the Energy Citations Database. SciTech Connect incorporates all of the R&D information from these two products into one search interface. It includes over 2.7 million citations, including citations to 1.5 million journal articles. SciTech Connect also has over 400,000 full-text DOE sponsored STI reports; most of these are post-1991, but over 140,000 of the reports were published prior to 1990.

<http://www.osti.gov/scitech/>

## 5. Particle Physics Journals and Conference Proceedings Series

- **CERN Journals List:** This list of journals and conference series publishing particle physics content provides information on Open Access, copyright policies and terms of use.

<http://library.web.cern.ch/oa/wherepublish>

- **INSPIRE Journals:** The database covers more than 3,400 journals publishing HEP-related articles.

<http://inspirehep.net/collection/journals>

## 6. Conference Databases

- **INSPIRE Conferences:** The database of more than 20,600 past, present and future conferences, schools, and meetings of interest to high-energy physics and related fields is searchable by title, acronym, series, date, location. Included are information about published proceedings, links to conference contributions in the INSPIRE HEP database, and links to the conference Web site when available. New conferences can be submitted from the entry page.

<http://inspirehep.net/conferences>

## 7. Research Institutions

- **INSPIRE Institutions:** The database of more than 10,800 institutes, laboratories, and university departments in which research on particle physics and astrophysics is performed covers six continents and over a hundred countries. Included are address and Web links where available as well as links to the papers from each institution in the HEP database, to scientists listed in HEPNames affiliated to this institution in the past or present and to experiments performed at this institution. Searches can be performed by name, acronym, location, etc. The site offers an alphabetical list by country as well as a list of the top 500 HEP and astrophysics institutions sorted by country.

<http://inspirehep.net/institutions>

## 8. People

- **INSPIRE HEPNames:** Searchable worldwide database of over 112,000 people associated with particle physics and related fields. The affiliation history of these researchers, their e-mail addresses, web pages, experiments they participated in, PhD advisor, information on their graduate students and links to their papers in the INSPIRE HEP, arXiv and ADS databases are provided as well as a user interface to update these informations.

<http://inspirehep.net/hepnames>

## 9. Experiments

- **INSPIRE Experiments:** Contains more than 2,700 past, present, and future experiments in particle physics. Lists both accelerator and non-accelerator experiments. Includes official experiment name and number, location, and collaboration lists. Simple searches by participant, title, experiment number, institution, date approved, accelerator, or detector, return a description of the experiment, including a complete list of authors, title, overview of the experiment's goals and methods, and a link to the experiment's web page if available. Publication lists distinguish articles in refereed journals, theses, technical or instrumentation papers and those which rank among Topcite at 50 or more citations.

<http://inspirehep.net/Experiments>

- **Cosmic ray/Gamma ray/Neutrino and similar experiments:** This extensive collection of experimental web sites is organized by focus of study and also by location. Additional sections link to educational materials, organizations, related Web sites, etc. The site is maintained at the Max Planck Institute for Nuclear Physics, Heidelberg.

<http://www.mpi-hd.mpg.de/hfm/CosmicRay/CosmicRaySites.html>

## 10. Jobs

- **AAS Job Register:** The American Astronomical Society publishes once a month graduate, postgraduate, faculty and other positions mainly in astronomy and astrophysics.

<http://jobregister.aas.org/>

- **APS Careers:** A gateway for physicists, students, and physics enthusiasts to information about physics jobs and careers. Physics job listings, career advice, upcoming workshops and meetings, and career and job related resources provided by the American Physical Society.

<http://www.aps.org/careers/employment>

- **brightrecruits.com:** A recruitment service run by IOP Publishing that connects employers from different industry sectors with jobseekers who have a background in physics and engineering.

<http://brightrecruits.com/>

- **IOP Careers:** Careers information and resources primarily aimed at university students are provided by the UK Institute of Physics.

<http://www.iop.org/careers/>

- **INSPIRE HEPJobs:** Lists academic and research jobs in high energy physics, nuclear physics, accelerator physics and astrophysics with the option to post a job or to receive email notices of new job listings. About 500 jobs are currently listed.

<http://inspirehep.net/jobs>

- **Physics Today Jobs:** Online recruitment advertising website for Physics Today magazine, published by the American Institute of Physics. Physics Today Jobs is the managing partner of the AIP Career Network, an online job board network for the physical science, engineering, and computing disciplines. 8,000 resumes are currently available, and more than 2,500 jobs were posted in 2012.

<http://www.physicstoday.org/jobs>

## 11. Software Repositories

### Particle Physics

- **FastJet:** FastJet is a software package for jet finding in pp and e+e- collisions. It includes fast native implementations of many sequential recombination clustering algorithms, plugins for access to a range of cone jet finders and tools for advanced jet manipulation.

<http://fastjet.fr/>

- **FermiTools:** Fermilab's software tools program provides a repository of Fermilab - developed software packages of value to the HEP community. Permits searching for packages by title or subject category.

<http://www.fnal.gov/fermitools/>

- **FreeHEP:** A collection of software and information about software useful in high-energy physics and adjacent disciplines, focusing on open-source software for data analysis and visualization. Searching can be done by title, subject, date acquired, date updated, or by browsing an alphabetical list of all packages.

<http://www.freehep.org/>

- **Geant4:** Geant4 is a toolkit for the simulation of the passage of particles through matter. Its areas of application include high energy, nuclear and accelerator physics, as well as studies in medical and space science.

<http://geant4.web.cern.ch/geant4/>

- **GenSer:** The Generator Services project collaborates with Monte Carlo (MC) generators authors and with LHC experiments in order to prepare validated LCG compliant code for both the theoretical and experimental communities at the LHC, sharing the user support duties, providing assistance for the development of the new object-oriented generators and guaranteeing the maintenance of the older packages on the LCG supported platforms. The project consists of the generators repository, validation, HepMC record and MCDB event databases.

<http://ph-dep-sft.web.cern.ch/project/generator-service-project-genser>

- **Hepforge:** A development environment for high-energy physics software development projects, in particular housing many event-generator related projects, that offers a ready-made, easy-to-use set of Web based tools, including shell account with up to date development tools, web page hosting, subversion and CVS code management systems, mailing lists, bug tracker and wiki system.

<http://www.hepforge.org/>

- **QUDA:** Library for performing calculations in lattice QCD on GPUs using NVIDIA's "C for CUDA" API. The current

release includes optimized solvers for Wilson, Clover-improved Wilson, Twisted mass, Improved staggered (asqtad or HISQ), Domain wall and Mobius fermion actions.

<http://lattice.github.com/quda/>

- **ROOT:** This framework for data processing in high-energy physics, born at CERN, offers applications to store, access, process, analyze and represent data or perform simulations.

<http://root.cern.ch>

- **tmLQCD:** This freely available software suite provides a set of tools to be used in lattice QCD simulations, mainly a HMC implementation for Wilson and Wilson twisted mass fermions and inverter for different versions of the Dirac operator.

<https://github.com/etmc/tmLQCD>

- **USQCD:** The software suite enables lattice QCD computations to be performed with high performance across a variety of architectures. The page contains links to the project web pages of the individual software modules, as well as to complete lattice QCD application packages which use them.

<http://usqcd-software.github.io>

### Astrophysics

- **ASCL:** The Astrophysics Source Code Library (ASCL) is a free online registry for source codes of interest to astronomers and astrophysicists and lists codes that have been used in research that has appeared in, or been submitted to, peer-reviewed publications.

<http://ascl.net>

- **Astropy:** The Astropy Project is a community effort to develop a single core package for Astronomy in Python and foster interoperability between Python astronomy packages

<http://www.astropy.org>

- **IRAF:** The Image Reduction and Analysis Facility is a general purpose software system for the reduction and analysis of astronomical data. IRAF is written and supported by the National Optical Astronomy Observatories (NOAO) in Tucson, Arizona.

<http://iraf.noao.edu/>

- **Starlink:** Starlink was a UK Project supporting astronomical data processing. It was shut down in 2005 but its open-source software continued to be developed at the Joint Astronomy Centre until March 2015. It is currently maintained by the East Asian Observatory. The open-source software products are a collection of applications and libraries, usually focused on a specific aspect of data reduction or analysis.

<http://starlink.eao.hawaii.edu/starlink>

- Links to a large number of astronomy software archives are listed at:

<http://heasarc.nasa.gov/docs/heasarc/astro-update/>

### Apps

- **arXiv mobile:** Android app for browsing and searching arXiv.org, and for reading, saving and sharing articles.

[play.google.com/store/apps/details?id=com.commonsware.android.arXiv](http://play.google.com/store/apps/details?id=com.commonsware.android.arXiv)

- **arXiv scanner:** Scans downloads folder for pdf files from arXiv. Adds title, authors and summary and makes all this information easily searchable from inside the application.

<https://play.google.com/store/apps/details?id=com.agio.arxiv.scanner>

- **aNarXiv:** arXiv viewer.

<http://github.com/nephoapp/anarxiv>

- **Collider:** This mobile app allows to see data from the ATLAS experiment at the LHC.

<http://collider.physics.ox.ac.uk/>

- **LHSee:** This smartphone app allows to see collisions from the Large Hadron Collider.

<http://www2.physics.ox.ac.uk/about-us/outreach/public/lhsee>

- **The Particles:** App for Apple iPad, Windows 8 and Microsoft Surface. Allows to browse a wealth of real event images and videos, read popular biographies of each of the particles and explore the A-Z of particle physics with its details and definitions of key concepts, laboratories and physicists. Developed by Science Photo Library in partnership with Prof. Frank Close.

<http://www.sciencephoto.com/apps/particles.html>

## 12. Data repositories

### Particle Physics

- **HepData:** The HepData Project, funded by the STFC (UK) and based at Durham University, has been built up over the past four decades as a unique repository for scattering data from experimental particle physics. It currently comprises the data points from plots and tables related to several thousand publications including those from the LHC.

<http://hepdata.cedar.ac.uk/>

The data from HEPData can also be accessed through INSPIRE. A new enhanced service, rebranded as HEPData, is under development and will be available at:

<http://hepdata.net>

- **CERN Open Data:** The CERN Open Data portal provides data from real collision events, produced by the experiments at the LHC; virtual machines to reproduce the analysis environment; and software to process them. It serves almost 30 TB of data and encourages use both for educational and research purposes.

<http://opendata.cern.ch>

- **ILDG:** The International Lattice Data Grid is an international organization which provides standards, services, methods and tools that facilitate the sharing and interchange of lattice QCD gauge configurations among scientific collaborations, by uniting their regional data grids. It offers semantic access with local tools to worldwide distributed data.

<http://www.usqcd.org/ildg/>

- **MCDB - Monte Carlo Database:** This central database of MC events aims to facilitate communication between Monte-Carlo experts and users of event samples in LHC collaborations. Having these events stored in a public place along with the corresponding documentation allows for direct cross checks of the performances on reference samples.

<http://mcdb.cern.ch/>

- **MCPLLOTS:** mcplots is a repository of Monte Carlo plots comparing High Energy Physics event generators to a wide variety of available experimental data. The site is supported by the LHC Physics Centre at CERN.

<http://mcplots.cern.ch/>

### Astrophysics

- **CfA Dataverse:** This astronomy data repository at Harvard is open to all scientific data from astronomical institutions worldwide.

<https://dataverse.harvard.edu/dataverse/cfa>

- **NASA's HEASARC:** The High Energy Astrophysics Science Archive Research Center (HEASARC) is the primary archive for NASA's (and other space agencies') missions dealing with electromagnetic radiation from extremely energetic phenomena ranging from black holes to the Big Bang.

<http://heasarc.gsfc.nasa.gov/>

- **LAMBDA @ HEASARC:** This data center for Cosmic Microwave Background research, a merger of the High Energy Astrophysics Science Archive Research Center (HEASARC) and the Legacy Archive for Microwave Background Data Analysis (LAMBDA),

provides archive data from NASA missions, software tools, and links to other sites of interest.

<http://lambda.gsfc.nasa.gov/>

- The NASA archives provide access to raw and processed datasets from numerous NASA missions.

Mikulski Archive for Space Telescopes (MAST): Hubble telescope, other missions (UV, optical):

<http://archive.stsci.edu/>

NASA/IPAC Infrared Science Archive: Spitzer, Herschel, Planck telescope, other missions:

<http://irsa.ipac.caltech.edu/>

- **NASA/IPAC Extragalactic Database (NED):** An astronomical database that collates and cross-correlates information on extragalactic objects. It contains their positions, basic data, and names as well as bibliographic references to published papers, and notes from catalogs and other publications. NED supports searches for objects and references, and offers browsing capabilities for abstracts of articles of extragalactic interest.

<http://ned.ipac.caltech.edu/>

- **SIMBAD:** The SIMBAD astronomical database provides basic data, cross-identifications, bibliography and measurements for astronomical objects outside the solar system. It can be queried by object name, coordinates and various criteria. Lists of objects and scripts can be submitted.

<http://simbad.u-strasbg.fr/simbad/>

- **Virtual Observatory:** The Virtual Observatory (VO) provides a suite of resources to query for original data from a large number of archives. Two main tools are provided. One runs queries across multiple databases (such as the SDSS database) and combines the results. The other queries hundreds of archives for all datasets that fall on a particular piece of sky.

<http://www.us-vo.org/>

### General Physics

- **NIST Physical Measurement Laboratory:** The National Institute of Standards and Technology provides access to physical reference data (physical constants, atomic spectroscopy data, x-ray and gamma-ray data, radiation dosimetry data, nuclear physics data and more) and measurements and calibrations data (dimensional and electromagnetic measurements). The site points to a general interest page, linking to exhibits of the Physical Measurement Laboratory in the NIST Virtual Museum.

<http://physics.nist.gov/>

- **Springer Materials - The Landolt-Börnstein Database:** Landolt-Börnstein is a data collection in all areas of physical sciences and engineering, among others particle physics, electronic structure and transport, magnetism, superconductivity. International experts scan the primary literature in more than 8,000 peer-reviewed journals and evaluate and select the most valid information to be included in the database. It includes more than 100,000 online documents, 1.2 million references, and covers 250,000 chemical substances. The search functionality is freely accessible and the search results are displayed in their context, whereas the full text is secured to subscribers.

<http://materials.springer.com/>

## 13. Data preservation activities

### Particle Physics

- **DASPOS:** A collective effort to explore the realisation of a viable data, software and computation preservation architecture in High Energy Physics

<https://daspos.crc.nd.edu>

- **DPHEP:** The efforts to define and coordinate Data Preservation and Long Term Analysis in HEP are coordinated by a study group formed to investigate the issues associated with these activities.

The group, DPHEP, was initiated during 2008-2009 and includes all HEP major experiments and labs.

Details of the organizational structure, the objectives, workshops and publications can be found on the website.

The group is endorsed by the International Committee for Future Accelerators (ICFA).

In July 2014 the DPHEP collaboration was formed as a result of the signature of the Collaboration Agreement by seven large funding agencies (others have since joined or are in the process of acquisition) and in June 2015 the first DPHEP Collaboration Workshop and Collaboration Board meeting took place.

<http://dphep.org>

## Astrophysics

More formal and advanced data preservation activity is ongoing in the field of Experimental Astrophysics, including:

- **SDSS** (Sloan Digital Sky Survey)  
<http://sdss.org>
- **Fermi Data**  
<http://fermi.gsfc.nasa.gov/ssc/data>
- **IVOA** (International Virtual Observatory Alliance)  
<http://www.ivoa.net/>

## 14. Particle Physics Education and Outreach Sites

### Science Educators' Networks:

- **IPPOG**: The International Particle Physics Outreach Group is a network of particle physicists, researchers, informal science educators and science explainers aiming to raise awareness, understanding and standards of global outreach efforts in particle physics and general science by providing discussion forums and regular information exchange for science institutions, proposing and implementing strategies to share lessons learned and best practices and promoting current outreach efforts of network members.  
<http://ippog.web.cern.ch>
- **Interactions.org**: Designed to serve as a central resource for communicators of particle physics. The daily updated site provides links to current particle physics news from the world's press, high-resolution photos and graphics from the particle physics laboratories of the world; links to education and outreach programs; information about science policy and funding; a glossary; and links to many educational sites.  
<http://www.interactions.org>
- **I2U2 (Interactions in Understanding the Universe)**: The I2U2 e-Labs use the Internet and distributed computing in high-school classes and provide an opportunity for students to organise and conduct authentic research; experience the environment of scientific collaborations; make real scientific contributions. It is supported by QuarkNet, NSF and DOE.  
<http://www.i2u2.org>

### Master Classes

- **CMS physics masterclass**: Lectures from active scientists give insight into methods of basic research, enabling the students to perform measurements on real data from the CMS experiment at the LHC. Like in an international research collaboration, the participants then discuss their results and compare with expectations.  
<http://cms.web.cern.ch/content/cms-physics-masterclass>
- **International Masterclasses**: Each year about 10000 high school students in 42 countries come to one of about 200 nearby universities or research centres for one day in order to unravel the mysteries of particle physics. Lectures from active scientists give insight in topics and methods of basic research at the fundaments of matter and

forces, enabling the students to perform measurements on real data from particle physics experiments themselves. At the end of each day, like in an international research collaboration, the participants join in a video conference for discussion and combination of their results.

<http://physicsmasterclasses.org/>

- **MINERVA**: MINERVA (Masterclass INVolving Event recognition visualised with Atlantis) is a masterclass tool for students to learn more about the ATLAS experiment at CERN, based on a simplified setup of the ATLAS event display, Atlantis.

<http://atlas-minerva.web.cern.ch/atlas-minerva/>

### General Sites

- **Contemporary Physics Education Project (CPEP)**: Provides charts, brochures, Web links, and classroom activities. Online interactive courses include: Fundamental Particles and Interactions; Plasma Physics and Fusion; History and Fate of the Universe; and Nuclear Science.

<http://www.cpepweb.org/>

### Particle Physics Lessons & Activities

- **Angels and Demons**: With the aim of looking at the myth versus the reality of antimatter and science at CERN this site offers teacher resources, slide shows and videos of talks given to teachers visiting CERN.

<http://angelsanddemons.web.cern.ch/>

- **ATLAS @ Home** A research project that uses volunteer computing to run simulations of the ATLAS experiment at CERN.

<http://atlasathome.cern.ch>

- **Big Bang Science: Exploring the origins of matter**: This Web site, produced by the Particle Physics and Astronomy Research Council of the UK (PPARC), explains what physicists are looking for with their giant instruments. It focuses on CERN particle detectors and on United Kingdom scientists' contribution to the search for the fundamental building blocks of matter.

<http://hepwww.rl.ac.uk/pub/bigbang/part1.html>

- **Cambridge Relativity and Cosmology**:

<http://www.damtp.cam.ac.uk/research/gr/public/index.html>

- **CAMELIA**: CAMELIA (Cross-platform Atlas Multimedia Educational Lab for Interactive Analysis) is a discovery tool for the general public, based on computer gaming technology.

<http://www.atlas.ch/camelia.html>

- **CERNland**: With a range of games, multimedia applications and films CERNland is a virtual theme park developed to bring the excitement of CERN's research to a young audience aged between 7 and 12. CERNland is designed to show children what is being done at CERN and inspire them with some physics at the same time.

<http://www.cernland.net/>

- **CollidingParticles**: A series of films following a team of physicists involved in research at the LHC.

<http://www.collidingparticles.com/>

- **Hands-On Universe**: This educational program enables students to investigate the Universe while applying tools and concepts from science, math and technology.

<http://handsonuniverse.org/>

- **Higgs Hunters**: A web-based citizen science project to help search for unknown exotic particles in the LHC data.

<http://HiggsHunters.org>

- **HYPATIA**: HYPATIA (Hybrid Pupil's Analysis Tool for Interactions in Atlas) is a tool for high school students to inspect the graphic visualization of products of particle collisions in the ATLAS detector at CERN.

<http://hypatia.phys.uoa.gr/>



- **Imagine the Universe:** This NASA site is intended for students age 14 and up and for anyone interested in learning about the universe.

<http://imagine.gsfc.nasa.gov/home.html>

- **In particular:** Podcast about physics and the process of discovering physics at the ATLAS experiment.

<https://itunes.apple.com/us/podcast/in-particular/id1001131655?mt=2>

- **Lancaster Particle Physics:** This site, suitable for 16+ students, offers a number of simulations and explanations of particle physics, including a section on the LHC.

<http://www.lppp.lancs.ac.uk/>

- **LHC @ home:** Volunteer computing platform to help physicists compare theory with experiment, in the search for new fundamental particles and answers to questions about the Universe.

<http://lhcathe.web.cern.ch>

The Test4Theory allows volunteers to run simulations of high-energy particle physics on their home computers. The results are submitted to a database which is used as a common resource by both experimental and theoretical scientists working on the Large Hadron Collider at CERN.

<http://lhcathe.web.cern.ch/projects/test4theory>

The SIXTRACK project allows users with Internet-connected computers to participate in advancing Accelerator Physics.

<http://lhcathe.web.cern.ch/projects/sixtrack>

- **Particle Adventure:** One of the most popular Web sites for learning the fundamentals of matter and force. An award-winning interactive tour of quarks, neutrinos, antimatter, extra dimensions, dark matter, accelerators and particle detectors from the Particle Data Group of Lawrence Berkeley National Laboratory. Simple elegant graphics and translations into 16 languages.

<http://particleadventure.org/>

- **Quarked! - Adventures in the Subatomic Universe:** This project, targeted to kids aged 7-12 (and their families), brings subatomic physics to life through a multimedia project including an interactive website, a facilitated program for museums and schools, and an educational outreach program.

<http://www.quarked.org/>

- **QuarkNet:** Brings the excitement of particle physics research to high school teachers and their students. Teachers join research groups at about 50 universities and labs across the country. These research groups are part of particle physics experiments at CERN or Fermilab. About 100,000 students from 500+ US high schools learn fundamental physics as they participate in inquiry-oriented investigations and analyze real data online. QuarkNet is supported in part by the National Science Foundation and the U.S. Department of Energy.

<https://quarknet.i2u2.org/>

- **Rewarding Learning videos about CERN:** The three videos based on interviews with scientists and engineers at CERN introduce pupils to CERN and the type of research and work undertaken there and are accompanied by teachers' notes.

<http://www.nicurriculum.org.uk/STEMWorks/resources/cern/index.asp>

#### Lab Education Offices

- **Argonne National Laboratory (ANL) Educational Programs:**

<http://www.anl.gov/education/>

- **Brookhaven National Laboratory (BNL) Educational Programs:** The Office of Educational Programs mission is to design, develop, implement, and facilitate workforce development and education initiatives that support the scientific mission at Brookhaven National Laboratory and the Department of Energy.

<http://www.bnl.gov/education/>

- **CERN:** The CERN education website offers informations about teacher programmes and educational resources for schools.

<http://education.web.cern.ch/education/>

- **DESY:** Offers courses for pupils and teachers as well as information for the general public, mostly in German.

[http://www.desy.de/information\\_services/education/](http://www.desy.de/information_services/education/)

- **FermiLab Education Office:** Provides education resources and information about activities for educators, physicists, students and visitors to the Lab. In addition to information on 25 programs, the site provides online data-based investigations for high school students, online versions of exhibits in the Lederman Science Center, links to particle physics discovery resources, web-based instructional resources, what works for education and outreach, and links to the Lederman Science Center and the Teacher Resource Center.

<http://ed.fnal.gov/>

- **Science Education at Jefferson Lab:**

<http://education.jlab.org/>

- **LBL Workforce Development and Education:** This group carries out Berkeley Labs mission to inspire and prepare the next generation of scientists, engineers, and technicians.

<http://csee.lbl.gov/>

#### Educational Programs of Experiments

- **ATLAS Discovery Quest:** One of several access points to ATLAS education and outreach pages. This page gives access to explanations of physical concepts, blogs, ATLAS facts, news, and information for students and teachers.

<http://www.atlas.ch/physics.html>

- **ATLAS eTours:** Give a description of the Large Hadron Collider, explain how the ATLAS detector at the LHC works and give an overview over the experiments and their physics goals.

<http://www.atlas.ch/etours.html>

- **Education and Outreach @ IceCube:**

<http://icecube.wisc.edu/outreach>

- **LIGO Science Education Center:** The LIGO (Laser Interferometer Gravitational-wave Observatory) Science Education Center has over 40 interactive, hands-on exhibits that relate to the science of LIGO. The site hosts field trips for students, teacher training programs, and tours for the general public. Visitors can explore science concepts such as light, gravity, waves, and interference; learn about LIGO's search for gravitational waves; and interact with scientists and engineers.

<https://ligo.caltech.edu/page/educational-resources>

- **Pierre Auger Observatory's Educational Pages:** The site offers information about cosmic rays and their detection, and provides material for students and teachers.

<https://www.auger.org/index.php/edu-outreach>

#### News

- **Asimmetrie:** Bimonthly magazine about particle physics published by INFN, the Istituto Nazionale di Fisica Nucleare

<http://www.asimmetrie.it/>

- **CERN Courier:**

<http://cerncourier.com/cws/latest/cern>

- **DESY inForm:**

[http://www.desy.de/aktuelles/desy\\_inform](http://www.desy.de/aktuelles/desy_inform)

- **Fermilab Today:**

<http://www.fnal.gov/pub/today/>

- **LC Newslite:** The newsletter of the Linear Collider community

<http://newslite.linearcollider.org/>

twitter: @ILCnewslite

- **IOP News:**  
<http://www.iop.org/news/>
- **JINR News:**  
[http://www1.jinr.ru/News/Jinrnews\\_index.html](http://www1.jinr.ru/News/Jinrnews_index.html)
- **News at Interactions.org:** The InterActions site provides news and press releases on particle physics.  
<http://www.interactions.org/cms/?pid=1000680>  
twitter: @particlenews
- **Symmetry:** This magazine about particle physics and its connections to other aspects of life and science, from interdisciplinary collaborations to policy to culture is published 6 times per year by Fermilab and SLAC.  
<http://www.symmetrymagazine.org/>  
twitter: @symmetrismag

#### Art in Physics

- **Arts@CERN:** The Collide@CERN residency programme brings together world-class artists and scientists in a free exchange of ideas.  
<http://arts.web.cern.ch/collide/> Accelerate@CERN is a country specific one-month research award for artists who have never spent time in a science lab before.  
<http://arts.web.cern.ch/acceleratecern>
- **Art of Physics Competition:** The Canadian Association of Physicists organizes this competition, the first was launched in 1992, with the aim of stimulating interest, especially among non-scientists, in some of the captivating imagery associated with physics. The challenge is to capture photographically a beautiful or unusual physics phenomenon and explain it in less than 200 words in terms that everyone can understand.  
<http://www.cap.ca/aop/art.html>

#### Blogs and Twitter

Lists of active blogs and tweets can be found on INSPIRE:

- **Scientist blogs:**  
<http://tinyurl.com/nmku27s>
- **Scientists with twitter accounts:**  
<http://tinyurl.com/nrg5k63>

- **Experiments with twitter accounts:**  
<http://tinyurl.com/q86kma8>
- **Institutions with twitter accounts:**  
<http://tinyurl.com/mzcm3nw>

List of physicists on Twitter at TrueSciPhi:

<http://truesciphi.org/phy.html>

Some selected particle physics related blogs:

- **ATLAS blog:**  
<http://www.atlas.ch/blog>
- **Physics arXiv blog:** MIT Technology Review blog on new ideas at arXiv.org.  
<http://www.technologyreview.com/blog/arxiv/>
- **Life and Physics:** Jon Butterworth's blog in the Guardian.  
<http://www.guardian.co.uk/science/life-and-physics>
- **Not Even Wrong:** Peter Woit's blog on topics in physics and mathematics.  
<http://www.math.columbia.edu/woit/wordpress/>
- **Preposterous Universe:** Theoretical physicist Sean Carroll's blog.  
<http://www.preposterousuniverse.com/>
- **Quantum diaries:** Thoughts on work and life from particle physicists from around the world.  
<http://www.quantumdiaries.org/>  
The US LHC blog gives a vivid account of the daily activity of US LHC researchers.  
<http://www.quantumdiaries.org/lab-81/>
- **Science blogs:** Launched in January 2006, ScienceBlogs features bloggers from a wide array of scientific disciplines, including physics.  
<http://scienceblogs.com/channel/physical-science/>
- **AstroBetter:** Blog with tips and tricks for professional astronomers  
<http://www.astrobetter.com/>

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\* There are also search limits in the Summary Tables for the Gauge and Higgs Bosons, the Leptons, the Quarks, and the Mesons.



## Gauge &amp; Higgs Boson Summary Table

## SUMMARY TABLES OF PARTICLE PROPERTIES

Extracted from the Particle Listings of the  
*Review of Particle Physics*

C. Patrignani *et al.* (Particle Data Group),  
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## GAUGE AND HIGGS BOSONS

 $\gamma$  (photon)

$$I(J^{PC}) = 0.1(1^{--})$$

Mass  $m < 1 \times 10^{-18}$  eV  
Charge  $q < 1 \times 10^{-35}$  e  
Mean life  $\tau$  = Stable

 $g$   
or gluon

$$I(J^P) = 0(1^-)$$

Mass  $m = 0$  [a]  
SU(3) color octet

## graviton

$$J = 2$$

Mass  $m < 6 \times 10^{-32}$  eV

## W

$$J = 1$$

Charge =  $\pm 1$  e  
Mass  $m = 80.385 \pm 0.015$  GeV  
 $W/Z$  mass ratio =  $0.88153 \pm 0.00017$   
 $m_Z - m_W = 10.803 \pm 0.015$  GeV  
 $m_{W^+} - m_{W^-} = -0.2 \pm 0.6$  GeV  
Full width  $\Gamma = 2.085 \pm 0.042$  GeV  
 $\langle N_{\pi^\pm} \rangle = 15.70 \pm 0.35$   
 $\langle N_{K^\pm} \rangle = 2.20 \pm 0.19$   
 $\langle N_p \rangle = 0.92 \pm 0.14$   
 $\langle N_{\text{charged}} \rangle = 19.39 \pm 0.08$

$W^-$  modes are charge conjugates of the modes below.

W <sup>+</sup> DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\frac{p}{(\text{MeV}/c)}$
$\ell^+ \nu$	[b] (10.86 $\pm$ 0.09) %	—	—
$e^+ \nu$	(10.71 $\pm$ 0.16) %	—	40192
$\mu^+ \nu$	(10.63 $\pm$ 0.15) %	—	40192
$\tau^+ \nu$	(11.38 $\pm$ 0.21) %	—	40173
hadrons	(67.41 $\pm$ 0.27) %	—	—
$\pi^+ \gamma$	< 7 $\times 10^{-6}$	95%	40192
$D_s^+ \gamma$	< 1.3 $\times 10^{-3}$	95%	40168
$c\bar{X}$	(33.3 $\pm$ 2.6) %	—	—
$c\bar{S}$	(31 $^{+13}_{-11}$ ) %	—	—
invisible	[c] (1.4 $\pm$ 2.9) %	—	—

## Z

$$J = 1$$

Charge = 0  
Mass  $m = 91.1876 \pm 0.0021$  GeV [d]  
Full width  $\Gamma = 2.4952 \pm 0.0023$  GeV  
 $\Gamma(\ell^+ \ell^-) = 83.984 \pm 0.086$  MeV [b]  
 $\Gamma(\text{invisible}) = 499.0 \pm 1.5$  MeV [e]  
 $\Gamma(\text{hadrons}) = 1744.4 \pm 2.0$  MeV  
 $\Gamma(\mu^+ \mu^-)/\Gamma(e^+ e^-) = 1.0009 \pm 0.0028$   
 $\Gamma(\tau^+ \tau^-)/\Gamma(e^+ e^-) = 1.0019 \pm 0.0032$  [f]

## Average charged multiplicity

$$\langle N_{\text{charged}} \rangle = 20.76 \pm 0.16 \quad (S = 2.1)$$

## Couplings to quarks and leptons

$g_V^\ell = -0.03783 \pm 0.00041$   
 $g_V^u = 0.25^{+0.07}_{-0.06}$   
 $g_V^d = -0.33^{+0.05}_{-0.06}$   
 $g_A^\ell = -0.50123 \pm 0.00026$   
 $g_A^u = 0.50^{+0.04}_{-0.06}$   
 $g_A^d = -0.523^{+0.050}_{-0.029}$   
 $g^{\nu\ell} = 0.5008 \pm 0.0008$   
 $g^{\nu e} = 0.53 \pm 0.09$   
 $g^{\nu\mu} = 0.502 \pm 0.017$

# Gauge & Higgs Boson Summary Table

## Asymmetry parameters <sup>[g]</sup>

$$\begin{aligned} A_e &= 0.1515 \pm 0.0019 \\ A_\mu &= 0.142 \pm 0.015 \\ A_\tau &= 0.143 \pm 0.004 \\ A_s &= 0.90 \pm 0.09 \\ A_c &= 0.670 \pm 0.027 \\ A_b &= 0.923 \pm 0.020 \end{aligned}$$

## Charge asymmetry (%) at Z pole

$$\begin{aligned} A_{FB}^{(0\ell)} &= 1.71 \pm 0.10 \\ A_{FB}^{(0u)} &= 4 \pm 7 \\ A_{FB}^{(0s)} &= 9.8 \pm 1.1 \\ A_{FB}^{(0c)} &= 7.07 \pm 0.35 \\ A_{FB}^{(0b)} &= 9.92 \pm 0.16 \end{aligned}$$

Z DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$e^+ e^-$	( 3.363 $\pm$ 0.004 ) %		45594
$\mu^+ \mu^-$	( 3.366 $\pm$ 0.007 ) %		45594
$\tau^+ \tau^-$	( 3.370 $\pm$ 0.008 ) %		45559
$\ell^+ \ell^-$	[b] ( 3.3658 $\pm$ 0.0023 ) %		—
$\ell^+ \ell^- \ell^+ \ell^-$	[h] ( 3.30 $\pm$ 0.31 ) $\times 10^{-6}$	S=1.1	45594
invisible	(20.00 $\pm$ 0.06 ) %		—
hadrons	(69.91 $\pm$ 0.06 ) %		—
( $u\bar{u} + c\bar{c}$ )/2	(11.6 $\pm$ 0.6 ) %		—
( $d\bar{d} + s\bar{s} + b\bar{b}$ )/3	(15.6 $\pm$ 0.4 ) %		—
$c\bar{c}$	(12.03 $\pm$ 0.21 ) %		—
$b\bar{b}$	(15.12 $\pm$ 0.05 ) %		—
$b\bar{b}b\bar{b}$	( 3.6 $\pm$ 1.3 ) $\times 10^{-4}$		—
$g g$	< 1.1 %	CL=95%	—
$\pi^0 \gamma$	< 2.01 $\times 10^{-5}$	CL=95%	45594
$\eta \gamma$	< 5.1 $\times 10^{-5}$	CL=95%	45592
$\omega \gamma$	< 6.5 $\times 10^{-4}$	CL=95%	45590
$\eta'(958) \gamma$	< 4.2 $\times 10^{-5}$	CL=95%	45589
$\gamma \gamma$	< 1.46 $\times 10^{-5}$	CL=95%	45594
$\pi^0 \pi^0$	< 1.52 $\times 10^{-5}$	CL=95%	45594
$\gamma \gamma \gamma$	< 1.0 $\times 10^{-5}$	CL=95%	45594
$\pi^\pm W^\mp$	[i] < 7 $\times 10^{-5}$	CL=95%	10162
$\rho^\pm W^\mp$	[i] < 8.3 $\times 10^{-5}$	CL=95%	10136
$J/\psi(1S) X$	( 3.51 $\pm$ 0.23 $\pm$ 0.25 ) $\times 10^{-3}$	S=1.1	—
$J/\psi(1S) \gamma$	< 2.6 $\times 10^{-6}$	CL=95%	45541
$\psi(2S) X$	( 1.60 $\pm$ 0.29 ) $\times 10^{-3}$		—
$\chi_{c1}(1P) X$	( 2.9 $\pm$ 0.7 ) $\times 10^{-3}$		—
$\chi_{c2}(1P) X$	< 3.2 $\times 10^{-3}$	CL=90%	—
$\Upsilon(1S) X + \Upsilon(2S) X$	( 1.0 $\pm$ 0.5 ) $\times 10^{-4}$		—
+ $\Upsilon(3S) X$			
$\Upsilon(1S) X$	< 3.4 $\times 10^{-6}$	CL=95%	—
$\Upsilon(2S) X$	< 6.5 $\times 10^{-6}$	CL=95%	—
$\Upsilon(3S) X$	< 5.4 $\times 10^{-6}$	CL=95%	—
$(D^0/\bar{D}^0) X$	(20.7 $\pm$ 2.0 ) %		—
$D^\pm X$	(12.2 $\pm$ 1.7 ) %		—
$D^*(2010)^\pm X$	[i] (11.4 $\pm$ 1.3 ) %		—
$D_{s1}(2536)^\pm X$	( 3.6 $\pm$ 0.8 ) $\times 10^{-3}$		—
$D_{sJ}(2573)^\pm X$	( 5.8 $\pm$ 2.2 ) $\times 10^{-3}$		—
$D^{*'}(2629)^\pm X$	searched for		—
$B^+ X$	[j] ( 6.08 $\pm$ 0.13 ) %		—
$B_s^0 X$	[j] ( 1.59 $\pm$ 0.13 ) %		—
$B_c^+ X$	searched for		—
$A_c^+ X$	( 1.54 $\pm$ 0.33 ) %		—
$\Xi_c^0 X$	seen		—
$\Xi_b X$	seen		—
$b$ -baryon X	[j] ( 1.38 $\pm$ 0.22 ) %		—
anomalous $\gamma$ + hadrons	[k] < 3.2 $\times 10^{-3}$	CL=95%	—
$e^+ e^- \gamma$	[k] < 5.2 $\times 10^{-4}$	CL=95%	45594
$\mu^+ \mu^- \gamma$	[k] < 5.6 $\times 10^{-4}$	CL=95%	45594
$\tau^+ \tau^- \gamma$	[k] < 7.3 $\times 10^{-4}$	CL=95%	45559
$\ell^+ \ell^- \gamma \gamma$	[l] < 6.8 $\times 10^{-6}$	CL=95%	—
$q\bar{q} \gamma \gamma$	[l] < 5.5 $\times 10^{-6}$	CL=95%	—
$\nu\bar{\nu} \gamma \gamma$	[l] < 3.1 $\times 10^{-6}$	CL=95%	45594
$e^\pm \mu^\mp$	LF [i] < 7.5 $\times 10^{-7}$	CL=95%	45594
$e^\pm \tau^\mp$	LF [i] < 9.8 $\times 10^{-6}$	CL=95%	45576
$\mu^\pm \tau^\mp$	LF [i] < 1.2 $\times 10^{-5}$	CL=95%	45576
$p e$	L,B < 1.8 $\times 10^{-6}$	CL=95%	45589
$p \mu$	L,B < 1.8 $\times 10^{-6}$	CL=95%	45589

## $H^0$

$$J = 0$$

$$\begin{aligned} \text{Mass } m &= 125.09 \pm 0.24 \text{ GeV} \\ \text{Full width } \Gamma &< 1.7 \text{ GeV, CL} = 95\% \end{aligned}$$

## $H^0$ Signal Strengths in Different Channels

See Listings for the latest unpublished results.

$$\begin{aligned} \text{Combined Final States} &= 1.10 \pm 0.11 \\ W W^* &= 1.08^{+0.18}_{-0.16} \\ Z Z^* &= 1.29^{+0.26}_{-0.23} \\ \gamma \gamma &= 1.16 \pm 0.18 \\ b\bar{b} &= 0.82 \pm 0.30 \quad (S = 1.1) \\ \mu^+ \mu^- &< 7.0, \text{ CL} = 95\% \\ \tau^+ \tau^- &= 1.12 \pm 0.23 \\ Z \gamma &< 9.5, \text{ CL} = 95\% \\ t\bar{t} H^0 \text{ Production} &= 2.3^{+0.7}_{-0.6} \end{aligned}$$

$H^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$e^+ e^-$	< 1.9 $\times 10^{-3}$	95%	62545
$J/\psi \gamma$	< 1.5 $\times 10^{-3}$	95%	62507
$\Upsilon(1S) \gamma$	< 1.3 $\times 10^{-3}$	95%	62187
$\Upsilon(2S) \gamma$	< 1.9 $\times 10^{-3}$	95%	62143
$\Upsilon(3S) \gamma$	< 1.3 $\times 10^{-3}$	95%	62116
$\mu \tau$	< 1.51 %	95%	62532
invisible	< 58 %	95%	—

## Neutral Higgs Bosons, Searches for

### Searches for a Higgs Boson with Standard Model Couplings

$$\text{Mass } m > 122 \text{ and none } 128\text{--}1000 \text{ GeV, CL} = 95\%$$

The limits for  $H_1^0$  and  $A^0$  in supersymmetric models refer to the  $m_h^{\text{max}}$  benchmark scenario for the supersymmetric parameters.

### $H_1^0$ in Supersymmetric Models ( $m_{H_1^0} < m_{H_2^0}$ )

$$\text{Mass } m > 92.8 \text{ GeV, CL} = 95\%$$

### $A^0$ Pseudoscalar Higgs Boson in Supersymmetric Models <sup>[n]</sup>

$$\text{Mass } m > 93.4 \text{ GeV, CL} = 95\% \quad \tan\beta > 0.4$$

## Charged Higgs Bosons ( $H^\pm$ and $H^{\pm\pm}$ ), Searches for

$$H^\pm \text{ Mass } m > 80 \text{ GeV, CL} = 95\%$$

## New Heavy Bosons ( $W'$ , $Z'$ , leptoquarks, etc.), Searches for

### Additional $W$ Bosons

$W'$  with standard couplings

$$\text{Mass } m > 3.710 \times 10^3 \text{ GeV, CL} = 95\% \quad (pp \text{ direct search})$$

$W_R$  (Right-handed  $W$  Boson)

$$\text{Mass } m > 715 \text{ GeV, CL} = 90\% \quad (\text{electroweak fit})$$

### Additional $Z$ Bosons

$Z_{\text{SM}}$  with standard couplings

$$\text{Mass } m > 2.900 \times 10^3 \text{ GeV, CL} = 95\% \quad (pp \text{ direct search})$$

$$\text{Mass } m > 1.500 \times 10^3 \text{ GeV, CL} = 95\% \quad (\text{electroweak fit})$$

$Z_{LR}$  of  $SU(2)_L \times SU(2)_R \times U(1)$  (with  $g_L = g_R$ )

$$\text{Mass } m > 630 \text{ GeV, CL} = 95\% \quad (p\bar{p} \text{ direct search})$$

$$\text{Mass } m > 1162 \text{ GeV, CL} = 95\% \quad (\text{electroweak fit})$$

$Z_\chi$  of  $SO(10) \rightarrow SU(5) \times U(1)_\chi$  (with  $g_\chi = e/\cos\theta_W$ )

$$\text{Mass } m > 2.620 \times 10^3 \text{ GeV, CL} = 95\% \quad (pp \text{ direct search})$$

$$\text{Mass } m > 1.141 \times 10^3 \text{ GeV, CL} = 95\% \quad (\text{electroweak fit})$$

$Z_\psi$  of  $E_6 \rightarrow SO(10) \times U(1)_\psi$  (with  $g_\psi = e/\cos\theta_W$ )

$$\text{Mass } m > 2.570 \times 10^3 \text{ GeV, CL} = 95\% \quad (pp \text{ direct search})$$

$$\text{Mass } m > 476 \text{ GeV, CL} = 95\% \quad (\text{electroweak fit})$$

$Z_\eta$  of  $E_6 \rightarrow SU(3) \times SU(2) \times U(1)_\eta$  (with  $g_\eta = e/\cos\theta_W$ )

$$\text{Mass } m > 1.870 \times 10^3 \text{ GeV, CL} = 95\% \quad (pp \text{ direct search})$$

$$\text{Mass } m > 619 \text{ GeV, CL} = 95\% \quad (\text{electroweak fit})$$

# Gauge & Higgs Boson Summary Table

## Scalar Leptoquarks

Mass  $m > 1050$  GeV, CL = 95% (1st generation, pair prod.)  
 Mass  $m > 304$  GeV, CL = 95% (1st generation, single prod.)  
 Mass  $m > 1000$  GeV, CL = 95% (2nd generation, pair prod.)  
 Mass  $m > 73$  GeV, CL = 95% (2nd generation, single prod.)  
 Mass  $m > 740$  GeV, CL = 95% (3rd generation, pair prod.)  
 (See the Particle Listings for assumptions on leptoquark quantum numbers and branching fractions.)

## Diquarks

Mass  $m > 4700$  GeV, CL = 95% ( $E_6$  diquark)

## Axiguon

Mass  $m > 3600$  GeV, CL = 95%

## Axions ( $A^0$ ) and Other Very Light Bosons, Searches for

The standard Peccei-Quinn axion is ruled out. Variants with reduced couplings or much smaller masses are constrained by various data. The Particle Listings in the full Review contain a Note discussing axion searches.

The best limit for the half-life of neutrinoless double beta decay with Majoron emission is  $> 7.2 \times 10^{24}$  years (CL = 90%).

## NOTES

In this Summary Table:

When a quantity has “(S = ...)” to its right, the error on the quantity has been enlarged by the “scale factor” S, defined as  $S = \sqrt{\chi^2/(N-1)}$ , where  $N$  is the number of measurements used in calculating the quantity. We do this when  $S > 1$ , which often indicates that the measurements are inconsistent. When  $S > 1.25$ , we also show in the Particle Listings an ideogram of the measurements. For more about S, see the Introduction.

A decay momentum  $p$  is given for each decay mode. For a 2-body decay,  $p$  is the momentum of each decay product in the rest frame of the decaying particle. For a 3-or-more-body decay,  $p$  is the largest momentum any of the products can have in this frame.

- [a] Theoretical value. A mass as large as a few MeV may not be precluded.
- [b]  $\ell$  indicates each type of lepton ( $e$ ,  $\mu$ , and  $\tau$ ), not sum over them.
- [c] This represents the width for the decay of the  $W$  boson into a charged particle with momentum below detectability,  $p < 200$  MeV.
- [d] The  $Z$ -boson mass listed here corresponds to a Breit-Wigner resonance parameter. It lies approximately 34 MeV above the real part of the position of the pole (in the energy-squared plane) in the  $Z$ -boson propagator.
- [e] This partial width takes into account  $Z$  decays into  $\nu\bar{\nu}$  and any other possible undetected modes.
- [f] This ratio has not been corrected for the  $\tau$  mass.
- [g] Here  $A \equiv 2g_V g_A / (g_V^2 + g_A^2)$ .
- [h] Here  $\ell$  indicates  $e$  or  $\mu$ .
- [i] The value is for the sum of the charge states or particle/antiparticle states indicated.
- [j] This value is updated using the product of (i) the  $Z \rightarrow b\bar{b}$  fraction from this listing and (ii) the  $b$ -hadron fraction in an unbiased sample of weakly decaying  $b$ -hadrons produced in  $Z$ -decays provided by the Heavy Flavor Averaging Group (HFAG, <http://www.slac.stanford.edu/xorg/hfag/osc/PDG2009/#FRACZ>).
- [k] See the  $Z$  Particle Listings for the  $\gamma$  energy range used in this measurement.
- [l] For  $m_{\gamma\gamma} = (60 \pm 5)$  GeV.
- [n] The limits assume no invisible decays.

# Lepton Summary Table

## LEPTONS

**e**

$$J = \frac{1}{2}$$

Mass  $m = (548.57990970 \pm 0.000000016) \times 10^{-6} \text{ u}$   
 Mass  $m = 0.5109989461 \pm 0.0000000031 \text{ MeV}$   
 $|m_{e^+} - m_{e^-}|/m < 8 \times 10^{-9}$ , CL = 90%  
 $|q_{e^+} + q_{e^-}|/e < 4 \times 10^{-8}$   
 Magnetic moment anomaly  
 $(g-2)/2 = (1159.65218091 \pm 0.00000026) \times 10^{-6}$   
 $(g_{e^+} - g_{e^-}) / g_{\text{average}} = (-0.5 \pm 2.1) \times 10^{-12}$   
 Electric dipole moment  $d < 0.87 \times 10^{-28} \text{ ecm}$ , CL = 90%  
 Mean life  $\tau > 6.6 \times 10^{28} \text{ yr}$ , CL = 90% [a]

 **$\mu$** 

$$J = \frac{1}{2}$$

Mass  $m = 0.1134289257 \pm 0.0000000025 \text{ u}$   
 Mass  $m = 105.6583745 \pm 0.00000024 \text{ MeV}$   
 Mean life  $\tau = (2.1969811 \pm 0.0000022) \times 10^{-6} \text{ s}$   
 $\tau_{\mu^+}/\tau_{\mu^-} = 1.00002 \pm 0.00008$   
 $c\tau = 658.6384 \text{ m}$   
 Magnetic moment anomaly  $(g-2)/2 = (11659209 \pm 6) \times 10^{-10}$   
 $(g_{\mu^+} - g_{\mu^-}) / g_{\text{average}} = (-0.11 \pm 0.12) \times 10^{-8}$   
 Electric dipole moment  $d = (-0.1 \pm 0.9) \times 10^{-19} \text{ ecm}$

### Decay parameters [b]

$\rho = 0.74979 \pm 0.00026$   
 $\eta = 0.057 \pm 0.034$   
 $\delta = 0.75047 \pm 0.00034$   
 $\xi P_{\mu} = 1.0009^{+0.0016}_{-0.0007} [c]$   
 $\xi P_{\mu} \delta / \rho = 1.0018^{+0.0016}_{-0.0007} [c]$   
 $\xi' = 1.00 \pm 0.04$   
 $\xi'' = 0.98 \pm 0.04$   
 $\alpha/A = (0 \pm 4) \times 10^{-3}$   
 $\alpha'/A = (-10 \pm 20) \times 10^{-3}$   
 $\beta/A = (4 \pm 6) \times 10^{-3}$   
 $\beta'/A = (2 \pm 7) \times 10^{-3}$   
 $\overline{\eta} = 0.02 \pm 0.08$

$\mu^+$  modes are charge conjugates of the modes below.

$\mu^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$P$ (MeV/c)
$e^- \overline{\nu}_e \nu_{\mu}$	$\approx 100\%$		53
$e^- \overline{\nu}_e \nu_{\mu} \gamma$	[d] $(1.4 \pm 0.4) \%$		53
$e^- \overline{\nu}_e \nu_{\mu} e^+ e^-$	[e] $(3.4 \pm 0.4) \times 10^{-5}$		53

### Lepton Family number (LF) violating modes

	LF	[f] < 1.2	%	90%	53
$e^- \nu_e \overline{\nu}_{\mu}$	LF	< 5.7	$\times 10^{-13}$	90%	53
$e^- \gamma$	LF	< 1.0	$\times 10^{-12}$	90%	53
$e^- e^+ e^-$	LF	< 7.2	$\times 10^{-11}$	90%	53

 **$\tau$** 

$$J = \frac{1}{2}$$

Mass  $m = 1776.86 \pm 0.12 \text{ MeV}$   
 $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}} < 2.8 \times 10^{-4}$ , CL = 90%  
 Mean life  $\tau = (290.3 \pm 0.5) \times 10^{-15} \text{ s}$   
 $c\tau = 87.03 \text{ } \mu\text{m}$   
 Magnetic moment anomaly  $> -0.052$  and  $< 0.013$ , CL = 95%  
 $\text{Re}(d_{\tau}) = -0.220$  to  $0.45 \times 10^{-16} \text{ ecm}$ , CL = 95%  
 $\text{Im}(d_{\tau}) = -0.250$  to  $0.0080 \times 10^{-16} \text{ ecm}$ , CL = 95%

### Weak dipole moment

$\text{Re}(d_{\tau}^W) < 0.50 \times 10^{-17} \text{ ecm}$ , CL = 95%  
 $\text{Im}(d_{\tau}^W) < 1.1 \times 10^{-17} \text{ ecm}$ , CL = 95%

### Weak anomalous magnetic dipole moment

$\text{Re}(\alpha_{\tau}^W) < 1.1 \times 10^{-3}$ , CL = 95%  
 $\text{Im}(\alpha_{\tau}^W) < 2.7 \times 10^{-3}$ , CL = 95%  
 $\tau^{\pm} \rightarrow \pi^{\pm} K_S^0 \nu_{\tau}$  (RATE DIFFERENCE) / (RATE SUM) =  
 $(-0.36 \pm 0.25)\%$

### Decay parameters

See the  $\tau$  Particle Listings for a note concerning  $\tau$ -decay parameters.

$\rho(e \text{ or } \mu) = 0.745 \pm 0.008$   
 $\rho(e) = 0.747 \pm 0.010$   
 $\rho(\mu) = 0.763 \pm 0.020$   
 $\xi(e \text{ or } \mu) = 0.985 \pm 0.030$   
 $\xi(e) = 0.994 \pm 0.040$   
 $\xi(\mu) = 1.030 \pm 0.059$   
 $\eta(e \text{ or } \mu) = 0.013 \pm 0.020$   
 $\eta(\mu) = 0.094 \pm 0.073$   
 $(\delta\xi)(e \text{ or } \mu) = 0.746 \pm 0.021$   
 $(\delta\xi)(e) = 0.734 \pm 0.028$   
 $(\delta\xi)(\mu) = 0.778 \pm 0.037$   
 $\xi(\pi) = 0.993 \pm 0.022$   
 $\xi(\rho) = 0.994 \pm 0.008$   
 $\xi(a_1) = 1.001 \pm 0.027$   
 $\xi(\text{all hadronic modes}) = 0.995 \pm 0.007$

$\tau^+$  modes are charge conjugates of the modes below. " $h^{\pm}$ " stands for  $\pi^{\pm}$  or  $K^{\pm}$ . " $e$ " stands for  $e$  or  $\mu$ . "Neutrals" stands for  $\gamma$ 's and/or  $\pi^0$ 's.

$\tau^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$P$ (MeV/c)
<b>Modes with one charged particle</b>			
particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_{\tau}$	(85.24 $\pm$ 0.06) %		—
("1-prong")			
particle $^- \geq 0$ neutrals $\geq 0 K_L^0 \nu_{\tau}$	(84.58 $\pm$ 0.06) %		—
$\mu^- \overline{\nu}_{\mu} \nu_{\tau}$	[g] (17.39 $\pm$ 0.04) %		885
$\mu^- \overline{\nu}_{\mu} \nu_{\tau} \gamma$	[e] (3.68 $\pm$ 0.10) $\times 10^{-3}$		885
$e^- \overline{\nu}_e \nu_{\tau}$	[g] (17.82 $\pm$ 0.04) %		888
$e^- \overline{\nu}_e \nu_{\tau} \gamma$	[e] (1.84 $\pm$ 0.05) %		888
$h^- \geq 0 K_L^0 \nu_{\tau}$	(12.03 $\pm$ 0.05) %		883
$h^- \nu_{\tau}$	(11.51 $\pm$ 0.05) %		883
$\pi^- \nu_{\tau}$	[g] (10.82 $\pm$ 0.05) %		883
$K^- \nu_{\tau}$	[g] (6.96 $\pm$ 0.10) $\times 10^{-3}$		820
$h^- \geq 1$ neutrals $\nu_{\tau}$	(37.00 $\pm$ 0.09) %		—
$h^- \geq 1 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	(36.51 $\pm$ 0.09) %		—
$h^- \pi^0 \nu_{\tau}$	(25.93 $\pm$ 0.09) %		878
$\pi^- \pi^0 \nu_{\tau}$	[g] (25.49 $\pm$ 0.09) %		878
$\pi^- \pi^0 \text{non-}\rho(770) \nu_{\tau}$	(3.0 $\pm$ 3.2) $\times 10^{-3}$		878
$K^- \pi^0 \nu_{\tau}$	[g] (4.33 $\pm$ 0.15) $\times 10^{-3}$		814
$h^- \geq 2 \pi^0 \nu_{\tau}$	(10.81 $\pm$ 0.09) %		—
$h^- 2 \pi^0 \nu_{\tau}$	(9.48 $\pm$ 0.10) %		862
$h^- 2 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	(9.32 $\pm$ 0.10) %		862
$\pi^- 2 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	[g] (9.26 $\pm$ 0.10) %		862
$\pi^- 2 \pi^0 \nu_{\tau} (\text{ex. } K^0),$	< 9	$\times 10^{-3}$ CL=95%	862
$\pi^- 2 \pi^0 \nu_{\tau} (\text{ex. } K^0),$	< 7	$\times 10^{-3}$ CL=95%	862
vector			
$K^- 2 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	[g] (6.5 $\pm$ 2.2) $\times 10^{-4}$		796
$h^- \geq 3 \pi^0 \nu_{\tau}$	(1.34 $\pm$ 0.07) %		—
$h^- \geq 3 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	(1.25 $\pm$ 0.07) %		—
$h^- 3 \pi^0 \nu_{\tau}$	(1.18 $\pm$ 0.07) %		836
$\pi^- 3 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	[g] (1.04 $\pm$ 0.07) %		836
$K^- 3 \pi^0 \nu_{\tau} (\text{ex. } K^0),$	[g] (4.8 $\pm$ 2.1) $\times 10^{-4}$		765
$\eta$			
$h^- 4 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	(1.6 $\pm$ 0.4) $\times 10^{-3}$		800
$h^- 4 \pi^0 \nu_{\tau} (\text{ex. } K^0, \eta)$	[g] (1.1 $\pm$ 0.4) $\times 10^{-3}$		800
$a_1(1260) \nu_{\tau} \rightarrow \pi^- \gamma \nu_{\tau}$	(3.8 $\pm$ 1.5) $\times 10^{-4}$		—
$K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_{\tau}$	(1.552 $\pm$ 0.029) %		820
$K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_{\tau}$	(8.59 $\pm$ 0.28) $\times 10^{-3}$		—
<b>Modes with <math>K^0</math>'s</b>			
$K_S^0 (\text{particles})^- \nu_{\tau}$	(9.44 $\pm$ 0.28) $\times 10^{-3}$		—
$h^- \overline{K}^0 \nu_{\tau}$	(9.87 $\pm$ 0.14) $\times 10^{-3}$		812
$\pi^- \overline{K}^0 \nu_{\tau}$	[g] (8.40 $\pm$ 0.14) $\times 10^{-3}$		812
$\pi^- \overline{K}^0$	(5.4 $\pm$ 2.1) $\times 10^{-4}$		812
(non- $K^*(892)^-$ ) $\nu_{\tau}$			
$K^- K^0 \nu_{\tau}$	[g] (1.48 $\pm$ 0.05) $\times 10^{-3}$		737
$K^- K^0 \geq 0 \pi^0 \nu_{\tau}$	(2.98 $\pm$ 0.08) $\times 10^{-3}$		737
$h^- \overline{K}^0 \pi^0 \nu_{\tau}$	(5.32 $\pm$ 0.13) $\times 10^{-3}$		794
$\pi^- \overline{K}^0 \pi^0 \nu_{\tau}$	[g] (3.82 $\pm$ 0.13) $\times 10^{-3}$		794
$\overline{K}^0 \rho^- \nu_{\tau}$	(2.2 $\pm$ 0.5) $\times 10^{-3}$		612
$K^- K^0 \pi^0 \nu_{\tau}$	[g] (1.50 $\pm$ 0.07) $\times 10^{-3}$		685
$\pi^- \overline{K}^0 \geq 1 \pi^0 \nu_{\tau}$	(4.08 $\pm$ 0.25) $\times 10^{-3}$		—
$\pi^- \overline{K}^0 \pi^0 \pi^0 \nu_{\tau} (\text{ex. } K^0)$	[g] (2.6 $\pm$ 2.3) $\times 10^{-4}$		763



## Lepton Summary Table

$K^- K^0 \pi^0 \pi^0 \nu_\tau$	$< 1.6$	$\times 10^{-4}$ CL=95%	619	$\pi^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$	$< 2.5$	$\times 10^{-3}$ CL=95%	794
$\pi^- K^0 \bar{K}^0 \nu_\tau$	$(1.55 \pm 0.24)$	$\times 10^{-3}$	682	$e^- e^- e^+ \bar{\nu}_e \nu_\tau$	$(2.8 \pm 1.5)$	$\times 10^{-5}$	888
$\pi^- K_S^0 K_S^0 \nu_\tau$	[g] $(2.33 \pm 0.07)$	$\times 10^{-4}$	682	$\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau$	$< 3.6$	$\times 10^{-5}$ CL=90%	885
$\pi^- K_S^0 K_L^0 \nu_\tau$	[g] $(1.08 \pm 0.24)$	$\times 10^{-3}$	682	<b>Modes with five charged particles</b>			
$\pi^- K_L^0 K_L^0 \nu_\tau$	$(2.33 \pm 0.07)$	$\times 10^{-4}$	682	$3h^- 2h^+ \geq 0$ neutrals $\nu_\tau$	$(9.9 \pm 0.4)$	$\times 10^{-4}$	794
$\pi^- K_S^0 \bar{K}^0 \pi^0 \nu_\tau$	$(3.6 \pm 1.2)$	$\times 10^{-4}$	614	(ex. $K_S^0 \rightarrow \pi^- \pi^+$ )			
$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	[g] $(1.82 \pm 0.21)$	$\times 10^{-5}$	614	("5-prong")			
$K^{*-} K^0 \pi^0 \nu_\tau \rightarrow$	$(1.08 \pm 0.21)$	$\times 10^{-5}$	-	$3h^- 2h^+ \nu_\tau$ (ex. $K^0$ )	$(8.22 \pm 0.32)$	$\times 10^{-4}$	794
$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$			-	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ )	$(8.21 \pm 0.31)$	$\times 10^{-4}$	794
$f_1(1285) \pi^- \nu_\tau \rightarrow$	$(6.8 \pm 1.5)$	$\times 10^{-6}$	-	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ , [g] $f_1(1285)$ )	$(7.69 \pm 0.30)$	$\times 10^{-4}$	-
$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$			-	$K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0$ )	[g] $(6 \pm 12)$	$\times 10^{-7}$	716
$f_1(1420) \pi^- \nu_\tau \rightarrow$	$(2.4 \pm 0.8)$	$\times 10^{-6}$	-	$K^+ 3\pi^- \pi^+ \nu_\tau$	$< 5.0$	$\times 10^{-6}$ CL=90%	716
$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	[g] $(3.2 \pm 1.2)$	$\times 10^{-4}$	614	$K^+ K^- 2\pi^- \pi^+ \nu_\tau$	$< 4.5$	$\times 10^{-7}$ CL=90%	528
$\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau$	$(1.82 \pm 0.21)$	$\times 10^{-5}$	614	$3h^- 2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.64 \pm 0.11)$	$\times 10^{-4}$	746
$K^- K_S^0 K_S^0 \nu_\tau$	$< 6.3$	$\times 10^{-7}$ CL=90%	466	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.62 \pm 0.11)$	$\times 10^{-4}$	746
$K^- K_S^0 K_S^0 \pi^0 \nu_\tau$	$< 4.0$	$\times 10^{-7}$ CL=90%	337	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta$ , $f_1(1285)$ )	$(1.11 \pm 0.10)$	$\times 10^{-4}$	-
$K^0 h^+ h^- h^- \geq 0$ neutrals $\nu_\tau$	$< 1.7$	$\times 10^{-3}$ CL=95%	760	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta$ , [g] $\omega, f_1(1285)$ )	$(3.8 \pm 0.9)$	$\times 10^{-5}$	-
$K^0 h^+ h^- h^- \nu_\tau$	[g] $(2.5 \pm 2.0)$	$\times 10^{-4}$	760	$K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	[g] $(1.1 \pm 0.6)$	$\times 10^{-6}$	657
<b>Modes with three charged particles</b>				$K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau$	$< 8$	$\times 10^{-7}$ CL=90%	657
$h^- h^- h^+ \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	$(15.21 \pm 0.06)$	%	861	$3h^- 2h^+ 2\pi^0 \nu_\tau$	$< 3.4$	$\times 10^{-6}$ CL=90%	687
$h^- h^- h^+ \geq 0$ neutrals $\nu_\tau$	$(14.55 \pm 0.06)$	%	861	<b>Miscellaneous other allowed modes</b>			
(ex. $K_S^0 \rightarrow \pi^+ \pi^-$ )				$(5\pi)^- \nu_\tau$	$(7.8 \pm 0.5)$	$\times 10^{-3}$	800
("3-prong")				$4h^- 3h^+ \geq 0$ neutrals $\nu_\tau$	$< 3.0$	$\times 10^{-7}$ CL=90%	682
$h^- h^- h^+ \nu_\tau$	$(9.80 \pm 0.05)$	%	861	("7-prong")			
$h^- h^- h^+ \nu_\tau$ (ex. $K^0$ )	$(9.46 \pm 0.05)$	%	861	$4h^- 3h^+ \pi^0$	$< 4.3$	$\times 10^{-7}$ CL=90%	682
$h^- h^- h^+ \nu_\tau$ (ex. $K^0, \omega$ )	$(9.43 \pm 0.05)$	%	861	$4h^- 3h^+ \pi^0 \nu_\tau$	$< 2.5$	$\times 10^{-7}$ CL=90%	612
$\pi^- \pi^+ \pi^- \nu_\tau$	$(9.31 \pm 0.05)$	%	861	$X^- (S=-1) \nu_\tau$	$(2.92 \pm 0.04)$	%	-
$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	$(9.02 \pm 0.05)$	%	861	$K^*(892)^- \geq 0$ neutrals $\geq$	$(1.42 \pm 0.18)$	%	S=1.4
$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ),	$< 2.4$	%	861	$0 K_L^0 \nu_\tau$			
non-axial vector				$K^*(892)^- \nu_\tau$	$(1.20 \pm 0.07)$	%	S=1.8
$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[g] $(8.99 \pm 0.05)$	%	861	$K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau$	$(7.83 \pm 0.26)$	$\times 10^{-3}$	-
$h^- h^- h^+ \geq 1$ neutrals $\nu_\tau$	$(5.29 \pm 0.05)$	%	-	$K^*(892)^0 K^- \geq 0$ neutrals $\nu_\tau$	$(3.2 \pm 1.4)$	$\times 10^{-3}$	542
$h^- h^- h^+ \geq 1 \pi^0 \nu_\tau$ (ex. $K^0$ )	$(5.09 \pm 0.05)$	%	-	$K^*(892)^0 K^- \nu_\tau$	$(2.1 \pm 0.4)$	$\times 10^{-3}$	542
$h^- h^- h^+ \pi^0 \nu_\tau$	$(4.76 \pm 0.05)$	%	834	$\bar{K}^*(892)^0 \pi^- \geq 0$ neutrals $\nu_\tau$	$(3.8 \pm 1.7)$	$\times 10^{-3}$	655
$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(4.57 \pm 0.05)$	%	834	$\bar{K}^*(892)^0 \pi^- \nu_\tau$	$(2.2 \pm 0.5)$	$\times 10^{-3}$	655
$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	$(2.79 \pm 0.07)$	%	834	$(\bar{K}^*(892) \pi)^- \nu_\tau \rightarrow$	$(1.0 \pm 0.4)$	$\times 10^{-3}$	-
$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$	$(4.62 \pm 0.05)$	%	834	$\pi^- \bar{K}^0 \pi^0 \nu_\tau$			
$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )	$(4.49 \pm 0.05)$	%	834	$K_1(1270)^- \nu_\tau$	$(4.7 \pm 1.1)$	$\times 10^{-3}$	433
$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	[g] $(2.74 \pm 0.07)$	%	834	$K_1(1400)^- \nu_\tau$	$(1.7 \pm 2.6)$	$\times 10^{-3}$	S=1.7
$h^- h^- h^+ \geq 2\pi^0 \nu_\tau$ (ex. $K^0$ )	$(5.17 \pm 0.31)$	$\times 10^{-3}$	-	$K^*(1410)^- \nu_\tau$	$(1.5 \pm 1.4)$	$\times 10^{-3}$	326
$h^- h^- h^+ 2\pi^0 \nu_\tau$	$(5.05 \pm 0.31)$	$\times 10^{-3}$	797	$K_0^*(1430)^- \nu_\tau$	$< 5$	$\times 10^{-4}$ CL=95%	317
$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0$ )	$(4.95 \pm 0.31)$	$\times 10^{-3}$	797	$K_2^*(1430)^- \nu_\tau$	$< 3$	$\times 10^{-3}$ CL=95%	317
$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[g] $(10 \pm 4)$	$\times 10^{-4}$	797	$\eta \pi^- \nu_\tau$	$< 9.9$	$\times 10^{-5}$ CL=95%	797
$h^- h^- h^+ 3\pi^0 \nu_\tau$	$(2.12 \pm 0.30)$	$\times 10^{-4}$	749	$\eta \pi^- \pi^0 \nu_\tau$	[g] $(1.39 \pm 0.07)$	$\times 10^{-3}$	778
$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.94 \pm 0.30)$	$\times 10^{-4}$	749	$\eta \pi^- \pi^0 \pi^0 \nu_\tau$	[g] $(1.9 \pm 0.4)$	$\times 10^{-4}$	746
$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta$ , $f_1(1285)$ )	$(1.7 \pm 0.4)$	$\times 10^{-4}$	-	$\eta K^- \nu_\tau$	[g] $(1.55 \pm 0.08)$	$\times 10^{-4}$	719
$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta$ , [g] $\omega, f_1(1285)$ )	$(1.4 \pm 2.7)$	$\times 10^{-5}$	-	$\eta K^*(892)^- \nu_\tau$	$(1.38 \pm 0.15)$	$\times 10^{-4}$	511
$K^- h^+ h^- \geq 0$ neutrals $\nu_\tau$	$(6.29 \pm 0.14)$	$\times 10^{-3}$	794	$\eta K^- \pi^0 \nu_\tau$	[g] $(4.8 \pm 1.2)$	$\times 10^{-5}$	665
$K^- h^+ \pi^- \nu_\tau$ (ex. $K^0$ )	$(4.37 \pm 0.07)$	$\times 10^{-3}$	794	$\eta K^- \pi^0$ (non- $K^*(892)$ ) $\nu_\tau$	$< 3.5$	$\times 10^{-5}$ CL=90%	-
$K^- h^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )	$(8.6 \pm 1.2)$	$\times 10^{-4}$	763	$\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau$	[g] $(9.4 \pm 1.5)$	$\times 10^{-5}$	661
$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$	$(4.77 \pm 0.14)$	$\times 10^{-3}$	794	$\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau$	$< 5.0$	$\times 10^{-5}$ CL=90%	590
$K^- \pi^+ \pi^- \geq$	$(3.73 \pm 0.13)$	$\times 10^{-3}$	794	$\eta K^- K^0 \nu_\tau$	$< 9.0$	$\times 10^{-6}$ CL=90%	430
$0\pi^0 \nu_\tau$ (ex. $K^0$ )				$\eta \pi^+ \pi^- \pi^- \geq 0$ neutrals $\nu_\tau$	$< 3$	$\times 10^{-3}$ CL=90%	744
$K^- \pi^+ \pi^- \nu_\tau$	$(3.45 \pm 0.07)$	$\times 10^{-3}$	794	$\eta \pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	[g] $(2.19 \pm 0.13)$	$\times 10^{-4}$	744
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	$(2.93 \pm 0.07)$	$\times 10^{-3}$	794	$\eta \pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, f_1(1285)$ )	$(9.9 \pm 1.6)$	$\times 10^{-5}$	-
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[g] $(2.93 \pm 0.07)$	$\times 10^{-3}$	794	$\eta a_1(1260)^- \nu_\tau \rightarrow \eta \pi^- \rho^0 \nu_\tau$	$< 3.9$	$\times 10^{-4}$ CL=90%	-
$K^- \rho^0 \nu_\tau \rightarrow$	$(1.4 \pm 0.5)$	$\times 10^{-3}$	-	$\eta \eta \pi^- \nu_\tau$	$< 7.4$	$\times 10^{-6}$ CL=90%	637
$K^- \pi^+ \pi^- \nu_\tau$				$\eta \eta \pi^- \pi^0 \nu_\tau$	$< 2.0$	$\times 10^{-4}$ CL=95%	559
$K^- \pi^+ \pi^- \pi^0 \nu_\tau$	$(1.31 \pm 0.12)$	$\times 10^{-3}$	763	$\eta \eta K^- \nu_\tau$	$< 3.0$	$\times 10^{-6}$ CL=90%	382
$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )	$(7.9 \pm 1.2)$	$\times 10^{-4}$	763	$\eta'(958) \pi^- \nu_\tau$	$< 4.0$	$\times 10^{-6}$ CL=90%	620
$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	$(7.6 \pm 1.2)$	$\times 10^{-4}$	763	$\eta'(958) \pi^- \pi^0 \nu_\tau$	$< 1.2$	$\times 10^{-5}$ CL=90%	591
$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	$(3.7 \pm 0.9)$	$\times 10^{-4}$	763	$\eta'(958) K^- \nu_\tau$	$< 2.4$	$\times 10^{-6}$ CL=90%	495
$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ ) [g]	$(3.9 \pm 1.4)$	$\times 10^{-4}$	763	$\phi \pi^- \nu_\tau$	$(3.4 \pm 0.6)$	$\times 10^{-5}$	585
$K^- \pi^+ K^- \geq 0$ neut. $\nu_\tau$	$< 9$	$\times 10^{-4}$ CL=95%	685	$\phi K^- \nu_\tau$	[g] $(4.4 \pm 1.6)$	$\times 10^{-5}$	445
$K^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$	$(1.496 \pm 0.033)$	$\times 10^{-3}$	685	$f_1(1285) \pi^- \nu_\tau$	$(3.9 \pm 0.5)$	$\times 10^{-4}$	S=1.9
$K^- K^+ \pi^- \nu_\tau$	[g] $(1.435 \pm 0.027)$	$\times 10^{-3}$	685	$f_1(1285) \pi^- \nu_\tau \rightarrow$	$(1.18 \pm 0.07)$	$\times 10^{-4}$	S=1.3
$K^- K^+ \pi^- \pi^0 \nu_\tau$	[g] $(6.1 \pm 1.8)$	$\times 10^{-5}$	618	$\eta \pi^- \pi^+ \pi^- \nu_\tau$			-
$K^- K^+ K^- \nu_\tau$	$(2.2 \pm 0.8)$	$\times 10^{-5}$	472	$f_1(1285) \pi^- \nu_\tau \rightarrow$	[g] $(5.2 \pm 0.4)$	$\times 10^{-5}$	-
$K^- K^+ K^- \nu_\tau$ (ex. $\phi$ )	$< 2.5$	$\times 10^{-6}$ CL=90%	-	$3\pi^- 2\pi^+ \nu_\tau$			-
$K^- K^+ K^- \pi^0 \nu_\tau$	$< 4.8$	$\times 10^{-6}$ CL=90%	345	$\pi(1300)^- \nu_\tau \rightarrow (\rho \pi)^- \nu_\tau \rightarrow$	$< 1.0$	$\times 10^{-4}$ CL=90%	-
				$(3\pi)^- \nu_\tau$			-

# Lepton Summary Table

$\pi(1300)^- \nu_\tau \rightarrow$	$< 1.9$	$\times 10^{-4}$ CL=90%	–
$((\pi\pi)_{S\text{-wave}} \pi)^- \nu_\tau \rightarrow$			
$(3\pi)^- \nu_\tau$			
$h^- \omega \geq 0 \text{ neutrals } \nu_\tau$	$(2.40 \pm 0.08) \%$		708
$h^- \omega \nu_\tau$	$(1.99 \pm 0.06) \%$		708
$\pi^- \omega \nu_\tau$	$[g] (1.95 \pm 0.06) \%$		708
$K^- \omega \nu_\tau$	$[g] (4.1 \pm 0.9) \times 10^{-4}$		610
$h^- \omega \pi^0 \nu_\tau$	$[g] (4.1 \pm 0.4) \times 10^{-3}$		684
$h^- \omega 2\pi^0 \nu_\tau$	$(1.4 \pm 0.5) \times 10^{-4}$		644
$\pi^- \omega 2\pi^0 \nu_\tau$	$[g] (7.1 \pm 1.6) \times 10^{-5}$		644
$h^- 2\omega \nu_\tau$	$< 5.4$	$\times 10^{-7}$ CL=90%	250
$2h^- h^+ \omega \nu_\tau$	$(1.20 \pm 0.22) \times 10^{-4}$		641
$2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0)$	$[g] (8.4 \pm 0.6) \times 10^{-5}$		641

## Lepton Family number (LF), Lepton number (L), or Baryon number (B) violating modes

L means lepton number violation (e.g.  $\tau^- \rightarrow e^+ \pi^- \pi^-$ ). Following common usage, LF means lepton family violation and not lepton number violation (e.g.  $\tau^- \rightarrow e^- \pi^+ \pi^-$ ). B means baryon number violation.

$e^- \gamma$	LF	$< 3.3$	$\times 10^{-8}$ CL=90%	888
$\mu^- \gamma$	LF	$< 4.4$	$\times 10^{-8}$ CL=90%	885
$e^- \pi^0$	LF	$< 8.0$	$\times 10^{-8}$ CL=90%	883
$\mu^- \pi^0$	LF	$< 1.1$	$\times 10^{-7}$ CL=90%	880
$e^- K_S^0$	LF	$< 2.6$	$\times 10^{-8}$ CL=90%	819
$\mu^- K_S^0$	LF	$< 2.3$	$\times 10^{-8}$ CL=90%	815
$e^- \eta$	LF	$< 9.2$	$\times 10^{-8}$ CL=90%	804
$\mu^- \eta$	LF	$< 6.5$	$\times 10^{-8}$ CL=90%	800
$e^- \rho^0$	LF	$< 1.8$	$\times 10^{-8}$ CL=90%	719
$\mu^- \rho^0$	LF	$< 1.2$	$\times 10^{-8}$ CL=90%	715
$e^- \omega$	LF	$< 4.8$	$\times 10^{-8}$ CL=90%	716
$\mu^- \omega$	LF	$< 4.7$	$\times 10^{-8}$ CL=90%	711
$e^- K^*(892)^0$	LF	$< 3.2$	$\times 10^{-8}$ CL=90%	665
$\mu^- K^*(892)^0$	LF	$< 5.9$	$\times 10^{-8}$ CL=90%	659
$e^- \bar{K}^*(892)^0$	LF	$< 3.4$	$\times 10^{-8}$ CL=90%	665
$\mu^- \bar{K}^*(892)^0$	LF	$< 7.0$	$\times 10^{-8}$ CL=90%	659
$e^- \eta'(958)$	LF	$< 1.6$	$\times 10^{-7}$ CL=90%	630
$\mu^- \eta'(958)$	LF	$< 1.3$	$\times 10^{-7}$ CL=90%	625
$e^- f_0(980) \rightarrow e^- \pi^+ \pi^-$	LF	$< 3.2$	$\times 10^{-8}$ CL=90%	–
$\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-$	LF	$< 3.4$	$\times 10^{-8}$ CL=90%	–
$e^- \phi$	LF	$< 3.1$	$\times 10^{-8}$ CL=90%	596
$\mu^- \phi$	LF	$< 8.4$	$\times 10^{-8}$ CL=90%	590
$e^- e^+ e^-$	LF	$< 2.7$	$\times 10^{-8}$ CL=90%	888
$e^- \mu^+ \mu^-$	LF	$< 2.7$	$\times 10^{-8}$ CL=90%	882
$e^+ \mu^- \mu^-$	LF	$< 1.7$	$\times 10^{-8}$ CL=90%	882
$\mu^- e^+ e^-$	LF	$< 1.8$	$\times 10^{-8}$ CL=90%	885
$\mu^+ e^- e^-$	LF	$< 1.5$	$\times 10^{-8}$ CL=90%	885
$\mu^- \mu^+ \mu^-$	LF	$< 2.1$	$\times 10^{-8}$ CL=90%	873
$e^- \pi^+ \pi^-$	LF	$< 2.3$	$\times 10^{-8}$ CL=90%	877
$e^+ \pi^- \pi^-$	L	$< 2.0$	$\times 10^{-8}$ CL=90%	877
$\mu^- \pi^+ \pi^-$	LF	$< 2.1$	$\times 10^{-8}$ CL=90%	866
$\mu^+ \pi^- \pi^-$	L	$< 3.9$	$\times 10^{-8}$ CL=90%	866
$e^- \pi^+ K^-$	LF	$< 3.7$	$\times 10^{-8}$ CL=90%	813
$e^- \pi^- K^+$	LF	$< 3.1$	$\times 10^{-8}$ CL=90%	813
$e^+ \pi^- K^-$	L	$< 3.2$	$\times 10^{-8}$ CL=90%	813
$e^- K_S^0 K_S^0$	LF	$< 7.1$	$\times 10^{-8}$ CL=90%	736
$e^- K^+ K^-$	LF	$< 3.4$	$\times 10^{-8}$ CL=90%	738
$e^+ K^- K^-$	L	$< 3.3$	$\times 10^{-8}$ CL=90%	738
$\mu^- \pi^+ K^-$	LF	$< 8.6$	$\times 10^{-8}$ CL=90%	800
$\mu^- \pi^- K^+$	LF	$< 4.5$	$\times 10^{-8}$ CL=90%	800
$\mu^+ \pi^- K^-$	L	$< 4.8$	$\times 10^{-8}$ CL=90%	800
$\mu^- K_S^0 K_S^0$	LF	$< 8.0$	$\times 10^{-8}$ CL=90%	696
$\mu^- K^+ K^-$	LF	$< 4.4$	$\times 10^{-8}$ CL=90%	699
$\mu^+ K^- K^-$	L	$< 4.7$	$\times 10^{-8}$ CL=90%	699
$e^- \pi^0 \pi^0$	LF	$< 6.5$	$\times 10^{-6}$ CL=90%	878
$\mu^- \pi^0 \pi^0$	LF	$< 1.4$	$\times 10^{-5}$ CL=90%	867
$e^- \eta \eta$	LF	$< 3.5$	$\times 10^{-5}$ CL=90%	699
$\mu^- \eta \eta$	LF	$< 6.0$	$\times 10^{-5}$ CL=90%	653
$e^- \pi^0 \eta$	LF	$< 2.4$	$\times 10^{-5}$ CL=90%	798
$\mu^- \pi^0 \eta$	LF	$< 2.2$	$\times 10^{-5}$ CL=90%	784
$p \mu^- \mu^-$	L,B	$< 4.4$	$\times 10^{-7}$ CL=90%	618
$\bar{p} \mu^+ \mu^-$	L,B	$< 3.3$	$\times 10^{-7}$ CL=90%	618
$\bar{p} \gamma$	L,B	$< 3.5$	$\times 10^{-6}$ CL=90%	641
$\bar{p} \pi^0$	L,B	$< 1.5$	$\times 10^{-5}$ CL=90%	632
$\bar{p} 2\pi^0$	L,B	$< 3.3$	$\times 10^{-5}$ CL=90%	604
$\bar{p} \eta$	L,B	$< 8.9$	$\times 10^{-6}$ CL=90%	475
$\bar{p} \pi^0 \eta$	L,B	$< 2.7$	$\times 10^{-5}$ CL=90%	360
$\Lambda \pi^-$	L,B	$< 7.2$	$\times 10^{-8}$ CL=90%	525

$\bar{\Lambda} \pi^-$	L,B	$< 1.4$	$\times 10^{-7}$ CL=90%	525
$e^- \text{light boson}$	LF	$< 2.7$	$\times 10^{-3}$ CL=95%	–
$\mu^- \text{light boson}$	LF	$< 5$	$\times 10^{-3}$ CL=95%	–

## Heavy Charged Lepton Searches

### $L^\pm$ – charged lepton

Mass  $m > 100.8$  GeV, CL = 95% <sup>[h]</sup> Decay to  $\nu W$ .

### $L^\pm$ – stable charged heavy lepton

Mass  $m > 102.6$  GeV, CL = 95%

## Neutrino Properties

See the note on “Neutrino properties listings” in the Particle Listings.

Mass  $m < 2$  eV (tritium decay)

Mean life/mass,  $\tau/m > 300$  s/eV, CL = 90% (reactor)

Mean life/mass,  $\tau/m > 7 \times 10^9$  s/eV (solar)

Mean life/mass,  $\tau/m > 15.4$  s/eV, CL = 90% (accelerator)

Magnetic moment  $\mu < 0.29 \times 10^{-10} \mu_B$ , CL = 90% (reactor)

## Number of Neutrino Types

Number  $N = 2.984 \pm 0.008$  (Standard Model fits to LEP-SLC data)

Number  $N = 2.92 \pm 0.05$  ( $S = 1.2$ ) (Direct measurement of invisible  $Z$  width)

## Neutrino Mixing

The following values are obtained through data analyses based on the 3-neutrino mixing scheme described in the review “Neutrino Mass, Mixing, and Oscillations” by K. Nakamura and S.T. Petcov in this *Review*.

$$\sin^2(\theta_{12}) = 0.304 \pm 0.014$$

$$\Delta m_{21}^2 = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2$$

$$\sin^2(\theta_{23}) = 0.51 \pm 0.05 \quad (\text{normal mass hierarchy})$$

$$\sin^2(\theta_{23}) = 0.50 \pm 0.05 \quad (\text{inverted mass hierarchy})$$

$$\Delta m_{32}^2 = (2.44 \pm 0.06) \times 10^{-3} \text{ eV}^2 \text{ [i]} \quad (\text{normal mass hierarchy})$$

$$\Delta m_{32}^2 = (2.51 \pm 0.06) \times 10^{-3} \text{ eV}^2 \text{ [i]} \quad (\text{inverted mass hierarchy})$$

$$\sin^2(\theta_{13}) = (2.19 \pm 0.12) \times 10^{-2}$$

### Stable Neutral Heavy Lepton Mass Limits

Mass  $m > 45.0$  GeV, CL = 95% (Dirac)

Mass  $m > 39.5$  GeV, CL = 95% (Majorana)

### Neutral Heavy Lepton Mass Limits

Mass  $m > 90.3$  GeV, CL = 95%

(Dirac  $\nu_L$  coupling to  $e, \mu, \tau$ ; conservative case( $\tau$ ))

Mass  $m > 80.5$  GeV, CL = 95%

(Majorana  $\nu_L$  coupling to  $e, \mu, \tau$ ; conservative case( $\tau$ ))

## NOTES

In this Summary Table:

When a quantity has “(S = ...)” to its right, the error on the quantity has been enlarged by the “scale factor” S, defined as  $S = \sqrt{\chi^2/(N-1)}$ , where N is the number of measurements used in calculating the quantity. We do this when  $S > 1$ , which often indicates that the measurements are inconsistent. When  $S > 1.25$ , we also show in the Particle Listings an ideogram of the measurements. For more about S, see the Introduction.

A decay momentum  $p$  is given for each decay mode. For a 2-body decay,  $p$  is the momentum of each decay product in the rest frame of the decaying particle. For a 3-or-more-body decay,  $p$  is the largest momentum any of the products can have in this frame.

# Lepton Summary Table

- [a] This is the best limit for the mode  $e^- \rightarrow \nu \gamma$ . The best limit for “electron disappearance” is  $6.4 \times 10^{24}$  yr.
- [b] See the “Note on Muon Decay Parameters” in the  $\mu$  Particle Listings for definitions and details.
- [c]  $P_\mu$  is the longitudinal polarization of the muon from pion decay. In standard  $V-A$  theory,  $P_\mu = 1$  and  $\rho = \delta = 3/4$ .
- [d] This only includes events with the  $\gamma$  energy  $> 10$  MeV. Since the  $e^- \bar{\nu}_e \nu_\mu$  and  $e^- \bar{\nu}_e \nu_\mu \gamma$  modes cannot be clearly separated, we regard the latter mode as a subset of the former.
- [e] See the relevant Particle Listings for the energy limits used in this measurement.
- [f] A test of additive vs. multiplicative lepton family number conservation.
- [g] Basis mode for the  $\tau$ .
- [h]  $L^\pm$  mass limit depends on decay assumptions; see the Full Listings.
- [i] The sign of  $\Delta m_{32}^2$  is not known at this time. The range quoted is for the absolute value.



## Meson Summary Table

# LIGHT UNFLAVORED MESONS ( $S = C = B = 0$ )

For  $I = 1$  ( $\pi, b, \rho, a$ ):  $u\bar{d}, (u\bar{u}-d\bar{d})/\sqrt{2}, d\bar{u}$ ;  
for  $I = 0$  ( $\eta, \eta', h, h', \omega, \phi, f, f'$ ):  $c_1(u\bar{u}+d\bar{d}) + c_2(s\bar{s})$

 $\pi^\pm$ 

$$I^G(J^P) = 1^-(0^-)$$

Mass  $m = 139.57018 \pm 0.00035$  MeV ( $S = 1.2$ )

Mean life  $\tau = (2.6033 \pm 0.0005) \times 10^{-8}$  s ( $S = 1.2$ )

$c\tau = 7.8045$  m

$\pi^\pm \rightarrow \ell^\pm \nu \gamma$  form factors <sup>[a]</sup>

$F_V = 0.0254 \pm 0.0017$

$F_A = 0.0119 \pm 0.0001$

$F_V$  slope parameter  $a = 0.10 \pm 0.06$

$R = 0.059^{+0.009}_{-0.008}$

$\pi^-$  modes are charge conjugates of the modes below.

For decay limits to particles which are not established, see the section on Searches for Axions and Other Very Light Bosons.

$\pi^\pm$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\mu^+ \nu_\mu$	[b] (99.98770 $\pm$ 0.00004) %		30
$\mu^+ \nu_\mu \gamma$	[c] ( 2.00 $\pm$ 0.25 ) $\times 10^{-4}$		30
$e^+ \nu_e$	[b] ( 1.230 $\pm$ 0.004 ) $\times 10^{-4}$		70
$e^+ \nu_e \gamma$	[c] ( 7.39 $\pm$ 0.05 ) $\times 10^{-7}$		70
$e^+ \nu_e \pi^0$	( 1.036 $\pm$ 0.006 ) $\times 10^{-8}$		4
$e^+ \nu_e e^+ e^-$	( 3.2 $\pm$ 0.5 ) $\times 10^{-9}$		70
$e^+ \nu_e \nu \bar{\nu}$	< 5 $\times 10^{-6}$	90%	70
<b>Lepton Family number (LF) or Lepton number (L) violating modes</b>			
$\mu^+ \bar{\nu}_e$	L [d] < 1.5 $\times 10^{-3}$	90%	30
$\mu^+ \nu_e$	LF [d] < 8.0 $\times 10^{-3}$	90%	30
$\mu^- e^+ e^+ \nu$	LF < 1.6 $\times 10^{-6}$	90%	30

 $\pi^0$ 

$$I^G(J^{PC}) = 1^-(0^{-+})$$

Mass  $m = 134.9766 \pm 0.0006$  MeV ( $S = 1.1$ )

$m_{\pi^\pm} - m_{\pi^0} = 4.5936 \pm 0.0005$  MeV

Mean life  $\tau = (8.52 \pm 0.18) \times 10^{-17}$  s ( $S = 1.2$ )

$c\tau = 25.5$  nm

For decay limits to particles which are not established, see the appropriate Search sections ( $A^0$  (axion) and Other Light Boson ( $X^0$ ) Searches, etc.).

$\pi^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$2\gamma$	(98.823 $\pm$ 0.034) %	S=1.5	67
$e^+ e^- \gamma$	( 1.174 $\pm$ 0.035 ) %	S=1.5	67
$\gamma$ positronium	( 1.82 $\pm$ 0.29 ) $\times 10^{-9}$		67
$e^+ e^+ e^- e^-$	( 3.34 $\pm$ 0.16 ) $\times 10^{-5}$		67
$e^+ e^-$	( 6.46 $\pm$ 0.33 ) $\times 10^{-8}$		67
$4\gamma$	< 2 $\times 10^{-8}$	CL=90%	67
$\nu \bar{\nu}$	[e] < 2.7 $\times 10^{-7}$	CL=90%	67
$\nu_e \bar{\nu}_e$	< 1.7 $\times 10^{-6}$	CL=90%	67
$\nu_\mu \bar{\nu}_\mu$	< 1.6 $\times 10^{-6}$	CL=90%	67
$\nu_\tau \bar{\nu}_\tau$	< 2.1 $\times 10^{-6}$	CL=90%	67
$\gamma \nu \bar{\nu}$	< 6 $\times 10^{-4}$	CL=90%	67
<b>Charge conjugation (C) or Lepton Family number (LF) violating modes</b>			
$3\gamma$	C < 3.1 $\times 10^{-8}$	CL=90%	67
$\mu^+ e^-$	LF < 3.8 $\times 10^{-10}$	CL=90%	26
$\mu^- e^+$	LF < 3.4 $\times 10^{-9}$	CL=90%	26
$\mu^+ e^- + \mu^- e^+$	LF < 3.6 $\times 10^{-10}$	CL=90%	26

 $\eta$ 

$$I^G(J^{PC}) = 0^+(0^{-+})$$

Mass  $m = 547.862 \pm 0.017$  MeV

Full width  $\Gamma = 1.31 \pm 0.05$  keV

## C-nonconserving decay parameters

$\pi^+ \pi^- \pi^0$  left-right asymmetry =  $(0.09 \pm^{0.11}_{0.12}) \times 10^{-2}$

$\pi^+ \pi^- \pi^0$  sextant asymmetry =  $(0.12 \pm^{0.10}_{0.11}) \times 10^{-2}$

$\pi^+ \pi^- \pi^0$  quadrant asymmetry =  $(-0.09 \pm 0.09) \times 10^{-2}$

$\pi^+ \pi^- \gamma$  left-right asymmetry =  $(0.9 \pm 0.4) \times 10^{-2}$

$\pi^+ \pi^- \gamma$   $\beta$  (D-wave) =  $-0.02 \pm 0.07$  ( $S = 1.3$ )

## CP-nonconserving decay parameters

$\pi^+ \pi^- e^+ e^-$  decay-plane asymmetry  $A_\phi = (-0.6 \pm 3.1) \times 10^{-2}$

## Dalitz plot parameter

$\pi^0 \pi^0 \pi^0$   $\alpha = -0.0318 \pm 0.0015$

$\eta$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
<b>Neutral modes</b>			
neutral modes	(72.12 $\pm$ 0.34) %	S=1.2	—
$2\gamma$	(39.41 $\pm$ 0.20) %	S=1.1	274
$3\pi^0$	(32.68 $\pm$ 0.23) %	S=1.1	179
$\pi^0 2\gamma$	( 2.56 $\pm$ 0.22 ) $\times 10^{-4}$		257
$2\pi^0 2\gamma$	< 1.2 $\times 10^{-3}$	CL=90%	238
$4\gamma$	< 2.8 $\times 10^{-4}$	CL=90%	274
invisible	< 1.0 $\times 10^{-4}$	CL=90%	—
<b>Charged modes</b>			
charged modes	(28.10 $\pm$ 0.34) %	S=1.2	—
$\pi^+ \pi^- \pi^0$	(22.92 $\pm$ 0.28) %	S=1.2	174
$\pi^+ \pi^- \gamma$	( 4.22 $\pm$ 0.08 ) %	S=1.1	236
$e^+ e^- \gamma$	( 6.9 $\pm$ 0.4 ) $\times 10^{-3}$	S=1.3	274
$\mu^+ \mu^- \gamma$	( 3.1 $\pm$ 0.4 ) $\times 10^{-4}$		253
$e^+ e^-$	< 2.3 $\times 10^{-6}$	CL=90%	274
$\mu^+ \mu^-$	( 5.8 $\pm$ 0.8 ) $\times 10^{-6}$		253
$2e^+ 2e^-$	( 2.40 $\pm$ 0.22 ) $\times 10^{-5}$		274
$\pi^+ \pi^- e^+ e^- (\gamma)$	( 2.68 $\pm$ 0.11 ) $\times 10^{-4}$		235
$e^+ e^- \mu^+ \mu^-$	< 1.6 $\times 10^{-4}$	CL=90%	253
$2\mu^+ 2\mu^-$	< 3.6 $\times 10^{-4}$	CL=90%	161
$\mu^+ \mu^- \pi^+ \pi^-$	< 3.6 $\times 10^{-4}$	CL=90%	113
$\pi^+ e^- \bar{\nu}_e + \text{c.c.}$	< 1.7 $\times 10^{-4}$	CL=90%	256
$\pi^+ \pi^- 2\gamma$	< 2.1 $\times 10^{-3}$		236
$\pi^+ \pi^- \pi^0 \gamma$	< 5 $\times 10^{-4}$	CL=90%	174
$\pi^0 \mu^+ \mu^- \gamma$	< 3 $\times 10^{-6}$	CL=90%	210

## Charge conjugation (C), Parity (P), Charge conjugation $\times$ Parity (CP), or Lepton Family number (LF) violating modes

$\pi^0 \gamma$	C	< 9 $\times 10^{-5}$	CL=90%	257
$\pi^+ \pi^-$	P, CP	< 1.3 $\times 10^{-5}$	CL=90%	236
$2\pi^0$	P, CP	< 3.5 $\times 10^{-4}$	CL=90%	238
$2\pi^0 \gamma$	C	< 5 $\times 10^{-4}$	CL=90%	238
$3\pi^0 \gamma$	C	< 6 $\times 10^{-5}$	CL=90%	179
$3\gamma$	C	< 1.6 $\times 10^{-5}$	CL=90%	274
$4\pi^0$	P, CP	< 6.9 $\times 10^{-7}$	CL=90%	40
$\pi^0 e^+ e^-$	C [f] < 4 $\times 10^{-5}$		CL=90%	257
$\pi^0 \mu^+ \mu^-$	C [f] < 5 $\times 10^{-6}$		CL=90%	210
$\mu^+ e^- + \mu^- e^+$	LF	< 6 $\times 10^{-6}$	CL=90%	264

$f_0(500)$  or  $\sigma$  <sup>[g]</sup>  
was  $f_0(600)$

$$I^G(J^{PC}) = 0^+(0^{++})$$

Mass  $m = (400-550)$  MeV

Full width  $\Gamma = (400-700)$  MeV

$f_0(500)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\pi \pi$	dominant	—
$\gamma \gamma$	seen	—

$\rho(770)$  <sup>[h]</sup>

$$I^G(J^{PC}) = 1^+(1^{--})$$

Mass  $m = 775.26 \pm 0.25$  MeV

Full width  $\Gamma = 149.1 \pm 0.8$  MeV

$\Gamma_{ee} = 7.04 \pm 0.06$  keV

# Meson Summary Table

<b><math>\rho(770)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$\pi^+\pi^-$	$\sim 100$	%	363
<b><math>\rho(770)^\pm</math> decays</b>			
$\pi^\pm\gamma$	( 4.5 $\pm$ 0.5 ) $\times 10^{-4}$	S=2.2	375
$\pi^\pm\eta$	< 6 $\times 10^{-3}$	CL=84%	152
$\pi^\pm\pi^+\pi^-\pi^0$	< 2.0 $\times 10^{-3}$	CL=84%	254
<b><math>\rho(770)^0</math> decays</b>			
$\pi^+\pi^-\gamma$	( 9.9 $\pm$ 1.6 ) $\times 10^{-3}$		362
$\pi^0\gamma$	( 6.0 $\pm$ 0.8 ) $\times 10^{-4}$		376
$\eta\gamma$	( 3.00 $\pm$ 0.20 ) $\times 10^{-4}$		194
$\pi^0\pi^0\gamma$	( 4.5 $\pm$ 0.8 ) $\times 10^{-5}$		363
$\mu^+\mu^-$	[i] ( 4.55 $\pm$ 0.28 ) $\times 10^{-5}$		373
$e^+e^-$	[i] ( 4.72 $\pm$ 0.05 ) $\times 10^{-5}$		388
$\pi^+\pi^-\pi^0$	( 1.01 $\pm$ $^{0.54}_{-0.36}$ ) $\times 10^{-4}$		323
$\pi^+\pi^-\pi^+\pi^-$	( 1.8 $\pm$ 0.9 ) $\times 10^{-5}$		251
$\pi^+\pi^-\pi^0\pi^0$	( 1.6 $\pm$ 0.8 ) $\times 10^{-5}$		257
$\pi^0e^+e^-$	< 1.2 $\times 10^{-5}$	CL=90%	376

$\omega(782)$

$$J^G(J^{PC}) = 0^-(1^{--})$$

Mass  $m = 782.65 \pm 0.12$  MeV (S = 1.9)  
Full width  $\Gamma = 8.49 \pm 0.08$  MeV  
 $\Gamma_{ee} = 0.60 \pm 0.02$  keV

<b><math>\omega(782)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$\pi^+\pi^-\pi^0$	(89.2 $\pm$ 0.7 ) %		327
$\pi^0\gamma$	( 8.28 $\pm$ 0.28 ) %	S=2.1	380
$\pi^+\pi^-$	( 1.53 $\pm$ $^{0.11}_{-0.13}$ ) %	S=1.2	366
neutrals (excluding $\pi^0\gamma$ )	( 8 $\pm$ $^{+8}_{-5}$ ) $\times 10^{-3}$	S=1.1	—
$\eta\gamma$	( 4.6 $\pm$ 0.4 ) $\times 10^{-4}$	S=1.1	200
$\pi^0e^+e^-$	( 7.7 $\pm$ 0.6 ) $\times 10^{-4}$		380
$\pi^0\mu^+\mu^-$	( 1.3 $\pm$ 0.4 ) $\times 10^{-4}$	S=2.1	349
$e^+e^-$	( 7.28 $\pm$ 0.14 ) $\times 10^{-5}$	S=1.3	391
$\pi^+\pi^-\pi^0\pi^0$	< 2 $\times 10^{-4}$	CL=90%	262
$\pi^+\pi^-\gamma$	< 3.6 $\times 10^{-3}$	CL=95%	366
$\pi^+\pi^-\pi^+\pi^-$	< 1 $\times 10^{-3}$	CL=90%	256
$\pi^0\pi^0\gamma$	( 6.6 $\pm$ 1.1 ) $\times 10^{-5}$		367
$\eta\pi^0\gamma$	< 3.3 $\times 10^{-5}$	CL=90%	162
$\mu^+\mu^-$	( 9.0 $\pm$ 3.1 ) $\times 10^{-5}$		377
$3\gamma$	< 1.9 $\times 10^{-4}$	CL=95%	391

Charge conjugation (C) violating modes

$\eta\pi^0$	C	< 2.1 $\times 10^{-4}$	CL=90%	162
$2\pi^0$	C	< 2.1 $\times 10^{-4}$	CL=90%	367
$3\pi^0$	C	< 2.3 $\times 10^{-4}$	CL=90%	330

$\eta'(958)$

$$J^G(J^{PC}) = 0^+(0^{-+})$$

Mass  $m = 957.78 \pm 0.06$  MeV  
Full width  $\Gamma = 0.197 \pm 0.009$  MeV

<b><math>\eta'(958)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\pi^+\pi^-\eta$	(42.9 $\pm$ 0.7 ) %		232
$\rho^0\gamma$ (including non-resonant $\pi^+\pi^-\gamma$ )	(29.1 $\pm$ 0.5 ) %		165
$\pi^0\pi^0\eta$	(22.3 $\pm$ 0.8 ) %		239
$\omega\gamma$	( 2.62 $\pm$ 0.13 ) %		159
$\omega e^+e^-$	( 2.0 $\pm$ 0.4 ) $\times 10^{-4}$		159
$\gamma\gamma$	( 2.21 $\pm$ 0.08 ) %		479
$3\pi^0$	( 2.20 $\pm$ 0.20 ) $\times 10^{-3}$		430
$\mu^+\mu^-\gamma$	( 1.08 $\pm$ 0.27 ) $\times 10^{-4}$		467
$\pi^+\pi^-\mu^+\mu^-$	< 2.9 $\times 10^{-5}$	90%	401
$\pi^+\pi^-\pi^0\pi^0$	( 3.82 $\pm$ 0.35 ) $\times 10^{-3}$		428
$\pi^0\rho^0$	< 4 %	90%	111
$2(\pi^+\pi^-)$	( 8.5 $\pm$ 0.9 ) $\times 10^{-5}$		372
$\pi^+\pi^-2\pi^0$	( 1.8 $\pm$ 0.4 ) $\times 10^{-4}$		376
$2(\pi^+\pi^-)$ neutrals	< 1 %	95%	—
$2(\pi^+\pi^-)\pi^0$	< 1.9 $\times 10^{-3}$	90%	298
$2(\pi^+\pi^-)2\pi^0$	< 1 %	95%	197

$3(\pi^+\pi^-)$	< 3.1 $\times 10^{-5}$	90%	189
$\pi^+\pi^-\pi^+e^-$	( 2.4 $\pm$ $^{+1.3}_{-1.0}$ ) $\times 10^{-3}$		458
$\pi^+e^-\nu_e + \text{c.c.}$	< 2.1 $\times 10^{-4}$	90%	469
$\gamma e^+e^-$	( 4.70 $\pm$ 0.30 ) $\times 10^{-4}$		479
$\pi^0\gamma\gamma$	< 8 $\times 10^{-4}$	90%	469
$4\pi^0$	< 3.2 $\times 10^{-4}$	90%	380
$e^+e^-$	< 5.6 $\times 10^{-9}$	90%	479
invisible	< 5 $\times 10^{-4}$	90%	—

Charge conjugation (C), Parity (P),  
Lepton family number (LF) violating modes

$\pi^+\pi^-$	$P, CP$	< 6 $\times 10^{-5}$	90%	458
$\pi^0\pi^0$	$P, CP$	< 4 $\times 10^{-4}$	90%	459
$\pi^0e^+e^-$	C	[f] < 1.4 $\times 10^{-3}$	90%	469
$\eta e^+e^-$	C	[f] < 2.4 $\times 10^{-3}$	90%	322
$3\gamma$	C	< 1.0 $\times 10^{-4}$	90%	479
$\mu^+\mu^-\pi^0$	C	[f] < 6.0 $\times 10^{-5}$	90%	445
$\mu^+\mu^-\eta$	C	[f] < 1.5 $\times 10^{-5}$	90%	273
$e\mu$	LF	< 4.7 $\times 10^{-4}$	90%	473

$f_0(980)$  [1]

$$J^G(J^{PC}) = 0^+(0^{++})$$

Mass  $m = 990 \pm 20$  MeV  
Full width  $\Gamma = 10$  to 100 MeV

<b><math>f_0(980)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\pi^+\pi^-$	dominant	476
$K\bar{K}$	seen	36
$\gamma\gamma$	seen	495

$a_0(980)$  [1]

$$J^G(J^{PC}) = 1^-(0^{++})$$

Mass  $m = 980 \pm 20$  MeV  
Full width  $\Gamma = 50$  to 100 MeV

<b><math>a_0(980)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\eta\pi$	dominant	319
$K\bar{K}$	seen	†
$\gamma\gamma$	seen	490

$\phi(1020)$

$$J^G(J^{PC}) = 0^-(1^{--})$$

Mass  $m = 1019.461 \pm 0.019$  MeV (S = 1.1)  
Full width  $\Gamma = 4.266 \pm 0.031$  MeV (S = 1.2)

<b><math>\phi(1020)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$K^+K^-$	(48.9 $\pm$ 0.5 ) %	S=1.1	127
$K_L^0K_S^0$	(34.2 $\pm$ 0.4 ) %	S=1.1	110
$\rho\pi + \pi^+\pi^-\pi^0$	(15.32 $\pm$ 0.32 ) %	S=1.1	—
$\eta\gamma$	( 1.309 $\pm$ 0.024 ) %	S=1.2	363
$\pi^0\gamma$	( 1.27 $\pm$ 0.06 ) $\times 10^{-3}$		501
$\ell^+\ell^-$	—		510
$e^+e^-$	( 2.954 $\pm$ 0.030 ) $\times 10^{-4}$	S=1.1	510
$\mu^+\mu^-$	( 2.87 $\pm$ 0.19 ) $\times 10^{-4}$		499
$\eta e^+e^-$	( 1.08 $\pm$ 0.04 ) $\times 10^{-4}$		363
$\pi^+\pi^-$	( 7.4 $\pm$ 1.3 ) $\times 10^{-5}$		490
$\omega\pi^0$	( 4.7 $\pm$ 0.5 ) $\times 10^{-5}$		172
$\omega\gamma$	< 5 %	CL=84%	209
$\rho\gamma$	< 1.2 $\times 10^{-5}$	CL=90%	215
$\pi^+\pi^-\gamma$	( 4.1 $\pm$ 1.3 ) $\times 10^{-5}$		490
$f_0(980)\gamma$	( 3.22 $\pm$ 0.19 ) $\times 10^{-4}$	S=1.1	29
$\pi^0\pi^0\gamma$	( 1.13 $\pm$ 0.06 ) $\times 10^{-4}$		492
$\pi^+\pi^-\pi^+\pi^-$	( 4.0 $\pm$ $^{+2.8}_{-2.2}$ ) $\times 10^{-6}$		410
$\pi^+\pi^+\pi^-\pi^-\pi^0$	< 4.6 $\times 10^{-6}$	CL=90%	342
$\pi^0e^+e^-$	( 1.12 $\pm$ 0.28 ) $\times 10^{-5}$		501
$\pi^0\eta\gamma$	( 7.27 $\pm$ 0.30 ) $\times 10^{-5}$	S=1.5	346
$a_0(980)\gamma$	( 7.6 $\pm$ 0.6 ) $\times 10^{-5}$		39
$K^0\bar{K}^0\gamma$	< 1.9 $\times 10^{-8}$	CL=90%	110
$\eta'(958)\gamma$	( 6.25 $\pm$ 0.21 ) $\times 10^{-5}$		60
$\eta\pi^0\pi^0\gamma$	< 2 $\times 10^{-5}$	CL=90%	293
$\mu^+\mu^-\gamma$	( 1.4 $\pm$ 0.5 ) $\times 10^{-5}$		499

## Meson Summary Table

$\rho\gamma\gamma$	< 1.2	$\times 10^{-4}$	CL=90%	215
$\eta\pi^+\pi^-$	< 1.8	$\times 10^{-5}$	CL=90%	288
$\eta\mu^+\mu^-$	< 9.4	$\times 10^{-6}$	CL=90%	321
$\eta U \rightarrow \eta e^+ e^-$	< 1	$\times 10^{-6}$	CL=90%	—

## Lepton Family number (LF) violating modes

$e^\pm\mu^\mp$	$LF$	< 2	$\times 10^{-6}$	CL=90%	504
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 **$h_1(1170)$**   $I^G(J^{PC}) = 0^-(1^+ -)$ 

Mass  $m = 1170 \pm 20$  MeV  
Full width  $\Gamma = 360 \pm 40$  MeV

$h_1(1170)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\rho\pi$	seen	308

 **$b_1(1235)$**   $I^G(J^{PC}) = 1^+(1^+ -)$ 

Mass  $m = 1229.5 \pm 3.2$  MeV (S = 1.6)  
Full width  $\Gamma = 142 \pm 9$  MeV (S = 1.2)

$b_1(1235)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$\omega\pi$	dominant		348
[D/S amplitude ratio = $0.277 \pm 0.027$ ]			
$\pi^\pm\gamma$	( $1.6 \pm 0.4$ ) $\times 10^{-3}$		607
$\eta\rho$	seen		†
$\pi^+\pi^+\pi^-\pi^0$	< 50 %	84%	535
$K^*(892)^\pm K^\mp$	seen		†
$(K\bar{K})^\pm\pi^0$	< 8 %	90%	248
$K_S^0 K_L^0\pi^\pm$	< 6 %	90%	235
$K_S^0 K_S^0\pi^\pm$	< 2 %	90%	235
$\phi\pi$	< 1.5 %	84%	147

 **$a_1(1260)^{[k]}$**   $I^G(J^{PC}) = 1^-(1^+ +)$ 

Mass  $m = 1230 \pm 40$  MeV <sup>[l]</sup>  
Full width  $\Gamma = 250$  to 600 MeV

$a_1(1260)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$(\rho\pi)_{S\text{-wave}}$	seen	353
$(\rho\pi)_{D\text{-wave}}$	seen	353
$(\rho(1450)\pi)_{S\text{-wave}}$	seen	†
$(\rho(1450)\pi)_{D\text{-wave}}$	seen	†
$\sigma\pi$	seen	—
$f_0(980)\pi$	not seen	179
$f_0(1370)\pi$	seen	†
$f_2(1270)\pi$	seen	†
$K\bar{K}^*(892) + \text{c.c.}$	seen	†
$\pi\gamma$	seen	608

 **$f_2(1270)$**   $I^G(J^{PC}) = 0^+(2^+ +)$ 

Mass  $m = 1275.5 \pm 0.8$  MeV  
Full width  $\Gamma = 186.7^{+2.2}_{-2.5}$  MeV (S = 1.4)

$f_2(1270)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$\rho$ (MeV/c)
$\pi\pi$	( $84.2^{+2.9}_{-0.9}$ ) %	S=1.1	623
$\pi^+\pi^-2\pi^0$	( $7.7^{+1.1}_{-3.2}$ ) %	S=1.2	563
$K\bar{K}$	( $4.6^{+0.5}_{-0.4}$ ) %	S=2.7	404
$2\pi^+2\pi^-$	( $2.8 \pm 0.4$ ) %	S=1.2	560
$\eta\eta$	( $4.0 \pm 0.8$ ) $\times 10^{-3}$	S=2.1	326
$4\pi^0$	( $3.0 \pm 1.0$ ) $\times 10^{-3}$		565
$\gamma\gamma$	( $1.42 \pm 0.24$ ) $\times 10^{-5}$	S=1.4	638
$\eta\pi\pi$	< 8 $\times 10^{-3}$	CL=95%	478
$K^0 K^- \pi^+ + \text{c.c.}$	< 3.4 $\times 10^{-3}$	CL=95%	293
$e^+ e^-$	< 6 $\times 10^{-10}$	CL=90%	638

 **$f_1(1285)$**   $I^G(J^{PC}) = 0^+(1^+ +)$ 

Mass  $m = 1282.0 \pm 0.5$  MeV (S = 1.8)  
Full width  $\Gamma = 24.1 \pm 1.0$  MeV (S = 1.3)

$f_1(1285)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$\rho$ (MeV/c)
$4\pi$	( $33.1^{+2.1}_{-1.8}$ ) %	S=1.3	568
$\pi^0\pi^0\pi^+\pi^-$	( $22.0^{+1.4}_{-1.2}$ ) %	S=1.3	566
$2\pi^+2\pi^-$	( $11.0^{+0.7}_{-0.6}$ ) %	S=1.3	563
$\rho^0\pi^+\pi^-$	( $11.0^{+0.7}_{-0.6}$ ) %	S=1.3	336
$\rho^0\rho^0$	seen		†
$4\pi^0$	< 7 $\times 10^{-4}$	CL=90%	568
$\eta\pi^+\pi^-$	( $35 \pm 15$ ) %		479
$\eta\pi\pi$	( $52.4^{+1.9}_{-2.2}$ ) %	S=1.2	482
$a_0(980)\pi$ [ignoring $a_0(980) \rightarrow K\bar{K}$ ]	( $36 \pm 7$ ) %		238
$\eta\pi\pi$ [excluding $a_0(980)\pi$ ]	( $16 \pm 7$ ) %		482
$K\bar{K}\pi$	( $9.0 \pm 0.4$ ) %	S=1.1	308
$K\bar{K}^*(892)$	not seen		†
$\pi^+\pi^-\pi^0$	( $3.0 \pm 0.9$ ) $\times 10^{-3}$		603
$\rho^\pm\pi^\mp$	< 3.1 $\times 10^{-3}$	CL=95%	390
$\gamma\rho^0$	( $5.5 \pm 1.3$ ) %	S=2.8	407
$\phi\gamma$	( $7.4 \pm 2.6$ ) $\times 10^{-4}$		236

 **$\eta(1295)$**   $I^G(J^{PC}) = 0^+(0^- +)$ 

Mass  $m = 1294 \pm 4$  MeV (S = 1.6)  
Full width  $\Gamma = 55 \pm 5$  MeV

$\eta(1295)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\eta\pi^+\pi^-$	seen	487
$a_0(980)\pi$	seen	248
$\eta\pi^0\pi^0$	seen	490
$\eta(\pi\pi)_{S\text{-wave}}$	seen	—

 **$\pi(1300)$**   $I^G(J^{PC}) = 1^-(0^- +)$ 

Mass  $m = 1300 \pm 100$  MeV <sup>[l]</sup>  
Full width  $\Gamma = 200$  to 600 MeV

$\pi(1300)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\rho\pi$	seen	404
$\pi(\pi\pi)_{S\text{-wave}}$	seen	—

 **$a_2(1320)$**   $I^G(J^{PC}) = 1^-(2^+ +)$ 

Mass  $m = 1318.3^{+0.5}_{-0.6}$  MeV (S = 1.2)  
Full width  $\Gamma = 107 \pm 5$  MeV <sup>[l]</sup>

$a_2(1320)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$\rho$ (MeV/c)
$3\pi$	( $70.1 \pm 2.7$ ) %	S=1.2	624
$\eta\pi\pi$	( $14.5 \pm 1.2$ ) %		535
$\omega\pi\pi$	( $10.6 \pm 3.2$ ) %	S=1.3	366
$K\bar{K}$	( $4.9 \pm 0.8$ ) %		437
$\eta'(958)\pi$	( $5.5 \pm 0.9$ ) $\times 10^{-3}$		288
$\pi^\pm\gamma$	( $2.91 \pm 0.27$ ) $\times 10^{-3}$		652
$\gamma\gamma$	( $9.4 \pm 0.7$ ) $\times 10^{-6}$		659
$e^+ e^-$	< 5 $\times 10^{-9}$	CL=90%	659

 **$f_0(1370)^{[j]}$**   $I^G(J^{PC}) = 0^+(0^+ +)$ 

Mass  $m = 1200$  to 1500 MeV  
Full width  $\Gamma = 200$  to 500 MeV

# Meson Summary Table

<b><math>f_0(1370)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\pi\pi$	seen	672
$4\pi$	seen	617
$4\pi^0$	seen	617
$2\pi^+2\pi^-$	seen	612
$\pi^+\pi^-2\pi^0$	seen	615
$\rho\rho$	dominant	†
$2(\pi\pi)_{S\text{-wave}}$	seen	–
$\pi(1300)\pi$	seen	†
$a_1(1260)\pi$	seen	35
$\eta\eta$	seen	411
$K\overline{K}$	seen	475
$K\overline{K}\eta\pi$	not seen	†
$6\pi$	not seen	508
$\omega\omega$	not seen	†
$\gamma\gamma$	seen	685
$e^+e^-$	not seen	685

$\pi_1(1400)$  [n]

$J^G(J^{PC}) = 1^-(1^--)$

Mass  $m = 1354 \pm 25$  MeV (S = 1.8)  
Full width  $\Gamma = 330 \pm 35$  MeV

<b><math>\pi_1(1400)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\eta\pi^0$	seen	557
$\eta\pi^-$	seen	556

$\eta(1405)$  [o]

$J^G(J^{PC}) = 0^+(0^{--})$

Mass  $m = 1408.8 \pm 1.8$  MeV [l] (S = 2.1)  
Full width  $\Gamma = 51.0 \pm 2.9$  MeV [l] (S = 1.8)

<b><math>\eta(1405)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$K\overline{K}\pi$	seen		424
$\eta\pi\pi$	seen		562
$a_0(980)\pi$	seen		345
$\eta(\pi\pi)_{S\text{-wave}}$	seen		–
$f_0(980)\eta$	seen		†
$4\pi$	seen		639
$\rho\rho$	<58 %	99.85%	†
$\rho^0\gamma$	seen		491
$K^*(892)K$	seen		123

$f_1(1420)$  [p]

$J^G(J^{PC}) = 0^+(1^{++})$

Mass  $m = 1426.4 \pm 0.9$  MeV (S = 1.1)  
Full width  $\Gamma = 54.9 \pm 2.6$  MeV

<b><math>f_1(1420)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\overline{K}\pi$	dominant	438
$K\overline{K}^*(892) + \text{c.c.}$	dominant	163
$\eta\pi\pi$	possibly seen	573
$\phi\gamma$	seen	349

$\omega(1420)$  [q]

$J^G(J^{PC}) = 0^-(1^{--})$

Mass  $m$  (1400–1450) MeV  
Full width  $\Gamma$  (180–250) MeV

<b><math>\omega(1420)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\rho\pi$	dominant	486
$\omega\pi\pi$	seen	444
$b_1(1235)\pi$	seen	125
$e^+e^-$	seen	710

$a_0(1450)$  [l]

$J^G(J^{PC}) = 1^-(0^{++})$

Mass  $m = 1474 \pm 19$  MeV  
Full width  $\Gamma = 265 \pm 13$  MeV

<b><math>a_0(1450)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\pi\eta$	seen	627
$\pi\eta'(958)$	seen	410
$K\overline{K}$	seen	547
$\omega\pi\pi$	seen	484
$a_0(980)\pi\pi$	seen	342
$\gamma\gamma$	seen	737

$\rho(1450)$  [r]

$J^G(J^{PC}) = 1^+(1^{--})$

Mass  $m = 1465 \pm 25$  MeV [l]  
Full width  $\Gamma = 400 \pm 60$  MeV [l]

<b><math>\rho(1450)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\pi\pi$	seen	720
$4\pi$	seen	669
$e^+e^-$	seen	732
$\eta\rho$	seen	311
$a_2(1320)\pi$	not seen	54
$K\overline{K}$	not seen	541
$K\overline{K}^*(892) + \text{c.c.}$	possibly seen	229
$\eta\gamma$	seen	630
$f_0(500)\gamma$	not seen	–
$f_0(980)\gamma$	not seen	398
$f_0(1370)\gamma$	not seen	92
$f_2(1270)\gamma$	not seen	177

$\eta(1475)$  [o]

$J^G(J^{PC}) = 0^+(0^{--})$

Mass  $m = 1476 \pm 4$  MeV (S = 1.3)  
Full width  $\Gamma = 85 \pm 9$  MeV (S = 1.5)

<b><math>\eta(1475)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\overline{K}\pi$	dominant	477
$K\overline{K}^*(892) + \text{c.c.}$	seen	245
$a_0(980)\pi$	seen	396
$\gamma\gamma$	seen	738
$K_S^0K_S^0\eta$	possibly seen	†

$f_0(1500)$  [n]

$J^G(J^{PC}) = 0^+(0^{++})$

Mass  $m = 1504 \pm 6$  MeV (S = 1.3)  
Full width  $\Gamma = 109 \pm 7$  MeV

<b><math>f_0(1500)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor	$\rho$ (MeV/c)
$\pi\pi$	(34.9±2.3) %	1.2	740
$\pi^+\pi^-$	seen		739
$2\pi^0$	seen		740
$4\pi$	(49.5±3.3) %	1.2	691
$4\pi^0$	seen		691
$2\pi^+2\pi^-$	seen		686
$2(\pi\pi)_{S\text{-wave}}$	seen		–
$\rho\rho$	seen		†
$\pi(1300)\pi$	seen		143
$a_1(1260)\pi$	seen		217
$\eta\eta$	( 5.1±0.9) %	1.4	515
$\eta\eta'(958)$	( 1.9±0.8) %	1.7	†
$K\overline{K}$	( 8.6±1.0) %	1.1	568
$\gamma\gamma$	not seen		752

$f_2'(1525)$

$J^G(J^{PC}) = 0^+(2^{++})$

Mass  $m = 1525 \pm 5$  MeV [l]  
Full width  $\Gamma = 73^{+6}_{-5}$  MeV [l]



Meson Summary Table

$f_2'(1525)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\overline{K}$	(88.7 $\pm$ 2.2 ) %	581
$\eta\eta$	(10.4 $\pm$ 2.2 ) %	530
$\pi\pi$	( 8.2 $\pm$ 1.5 ) $\times 10^{-3}$	750
$\gamma\gamma$	( 1.10 $\pm$ 0.14) $\times 10^{-6}$	763

$\pi_1(1600)$  <sup>[n]</sup>

$I^G(J^{PC}) = 1^-(1^--)$

Mass  $m = 1662^{+8}_{-9}$  MeV

Full width  $\Gamma = 241 \pm 40$  MeV    (S = 1.4)

$\pi_1(1600)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\pi\pi\pi$	not seen	803
$\rho^0\pi^-$	not seen	641
$f_2(1270)\pi^-$	not seen	318
$b_1(1235)\pi$	seen	357
$\eta'(958)\pi^-$	seen	543
$f_1(1285)\pi$	seen	314

$\eta_2(1645)$

$I^G(J^{PC}) = 0^+(2^--)$

Mass  $m = 1617 \pm 5$  MeV

Full width  $\Gamma = 181 \pm 11$  MeV

$\eta_2(1645)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$a_2(1320)\pi$	seen	242
$K\overline{K}\pi$	seen	580
$K^*\overline{K}$	seen	404
$\eta\pi^+\pi^-$	seen	685
$a_0(980)\pi$	seen	499
$f_2(1270)\eta$	not seen	†

$\omega(1650)$  <sup>[s]</sup>

$I^G(J^{PC}) = 0^-(1^--)$

Mass  $m = 1670 \pm 30$  MeV

Full width  $\Gamma = 315 \pm 35$  MeV

$\omega(1650)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\rho\pi$	seen	647
$\omega\pi\pi$	seen	617
$\omega\eta$	seen	500
$e^+e^-$	seen	835

$\omega_3(1670)$

$I^G(J^{PC}) = 0^-(3^--)$

Mass  $m = 1667 \pm 4$  MeV

Full width  $\Gamma = 168 \pm 10$  MeV <sup>[l]</sup>

$\omega_3(1670)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\rho\pi$	seen	645
$\omega\pi\pi$	seen	615
$b_1(1235)\pi$	possibly seen	361

$\pi_2(1670)$

$I^G(J^{PC}) = 1^-(2^--)$

Mass  $m = 1672.2 \pm 3.0$  MeV <sup>[l]</sup>    (S = 1.4)

Full width  $\Gamma = 260 \pm 9$  MeV <sup>[l]</sup>    (S = 1.2)

$\pi_2(1670)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$3\pi$	(95.8 $\pm$ 1.4) %		809
$f_2(1270)\pi$	(56.3 $\pm$ 3.2) %		328
$\rho\pi$	(31 $\pm$ 4 ) %		648
$\sigma\pi$	(10.9 $\pm$ 3.4) %		–
$\pi(\pi\pi)s$ -wave	( 8.7 $\pm$ 3.4) %		–
$K\overline{K}^*(892) + c.c.$	( 4.2 $\pm$ 1.4) %		455
$\omega\rho$	( 2.7 $\pm$ 1.1) %		304
$\pi^\pm\gamma$	( 7.0 $\pm$ 1.1) $\times 10^{-4}$		830
$\gamma\gamma$	< 2.8 $\times 10^{-7}$	90%	836
$\rho(1450)\pi$	< 3.6 $\times 10^{-3}$	97.7%	147
$b_1(1235)\pi$	< 1.9 $\times 10^{-3}$	97.7%	365
$f_1(1285)\pi$	possibly seen		323
$a_2(1320)\pi$	not seen		292

$\phi(1680)$

$I^G(J^{PC}) = 0^-(1^--)$

Mass  $m = 1680 \pm 20$  MeV <sup>[l]</sup>

Full width  $\Gamma = 150 \pm 50$  MeV <sup>[l]</sup>

$\phi(1680)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\overline{K}^*(892) + c.c.$	dominant	462
$K_S^0 K\pi$	seen	621
$K\overline{K}$	seen	680
$e^+e^-$	seen	840
$\omega\pi\pi$	not seen	623
$K^+K^-\pi^+\pi^-$	seen	544
$\eta\phi$	seen	290
$\eta\gamma$	seen	751

$\rho_3(1690)$

$I^G(J^{PC}) = 1^+(3^--)$

Mass  $m = 1688.8 \pm 2.1$  MeV <sup>[l]</sup>

Full width  $\Gamma = 161 \pm 10$  MeV <sup>[l]</sup>    (S = 1.5)

$\rho_3(1690)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor	$\rho$ (MeV/c)
$4\pi$	(71.1 $\pm$ 1.9 ) %		790
$\pi^\pm\pi^+\pi^-\pi^0$	(67 $\pm$ 22 ) %		787
$\omega\pi$	(16 $\pm$ 6 ) %		655
$\pi\pi$	(23.6 $\pm$ 1.3 ) %		834
$K\overline{K}\pi$	( 3.8 $\pm$ 1.2 ) %		629
$K\overline{K}$	( 1.58 $\pm$ 0.26) %	1.2	685
$\eta\pi^+\pi^-$	seen		727
$\rho(770)\eta$	seen		520
$\pi\pi\rho$	seen		633
Excluding $2\rho$ and $a_2(1320)\pi$ .			
$a_2(1320)\pi$	seen		307
$\rho\rho$	seen		335

$\rho(1700)$  <sup>[r]</sup>

$I^G(J^{PC}) = 1^+(1^--)$

Mass  $m = 1720 \pm 20$  MeV <sup>[l]</sup>    ( $\eta\rho^0$  and  $\pi^+\pi^-$  modes)

Full width  $\Gamma = 250 \pm 100$  MeV <sup>[l]</sup>    ( $\eta\rho^0$  and  $\pi^+\pi^-$  modes)

$\rho(1700)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$2(\pi^+\pi^-)$	large	803
$\rho\pi\pi$	dominant	653
$\rho^0\pi^+\pi^-$	large	651
$\rho^\pm\pi^\mp\pi^0$	large	652
$a_1(1260)\pi$	seen	404
$h_1(1170)\pi$	seen	447
$\pi(1300)\pi$	seen	349
$\rho\rho$	seen	372
$\pi^+\pi^-$	seen	849
$\pi\pi$	seen	849
$K\overline{K}^*(892) + c.c.$	seen	496
$\eta\rho$	seen	545
$a_2(1320)\pi$	not seen	334
$K\overline{K}$	seen	704
$e^+e^-$	seen	860
$\pi^0\omega$	seen	674

# Meson Summary Table

<b><math>f_0(1710)</math></b> <sup>[†]</sup>	$I^G(J^{PC}) = 0^+(0^+ +)$	
Mass $m = 1723^{+6}_{-5}$ MeV    (S = 1.6)		
Full width $\Gamma = 139 \pm 8$ MeV    (S = 1.1)		
<b><math>f_0(1710)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\bar{K}$	seen	706
$\eta\eta$	seen	665
$\pi\pi$	seen	851
$\omega\omega$	seen	360

<b><math>\pi(1800)</math></b>	$I^G(J^{PC}) = 1^-(0^- +)$	
Mass $m = 1812 \pm 12$ MeV    (S = 2.3)		
Full width $\Gamma = 208 \pm 12$ MeV		
<b><math>\pi(1800)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\pi^+\pi^-\pi^-$	seen	879
$f_0(500)\pi^-$	seen	—
$f_0(980)\pi^-$	seen	625
$f_0(1370)\pi^-$	seen	368
$f_0(1500)\pi^-$	not seen	250
$\rho\pi^-$	not seen	732
$\eta\eta\pi^-$	seen	661
$a_0(980)\eta$	seen	473
$a_2(1320)\eta$	not seen	†
$f_2(1270)\pi$	not seen	442
$f_0(1370)\pi^-$	not seen	368
$f_0(1500)\pi^-$	seen	250
$\eta\eta'(958)\pi^-$	seen	375
$K_0^*(1430)K^-$	seen	†
$K^*(892)K^-$	not seen	570

<b><math>\phi_3(1850)</math></b>	$I^G(J^{PC}) = 0^-(3^{--})$	
Mass $m = 1854 \pm 7$ MeV		
Full width $\Gamma = 87^{+28}_{-23}$ MeV    (S = 1.2)		
<b><math>\phi_3(1850)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\bar{K}$	seen	785
$K\bar{K}^*(892) + \text{c.c.}$	seen	602

<b><math>\pi_2(1880)</math></b>	$I^G(J^{PC}) = 1^-(2^- +)$	
Mass $m = 1895 \pm 16$ MeV		
Full width $\Gamma = 235 \pm 34$ MeV		

<b><math>f_2(1950)</math></b>	$I^G(J^{PC}) = 0^+(2^+ +)$	
Mass $m = 1944 \pm 12$ MeV    (S = 1.5)		
Full width $\Gamma = 472 \pm 18$ MeV		
<b><math>f_2(1950)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K^*(892)\bar{K}^*(892)$	seen	387
$\pi^+\pi^-$	seen	962
$\pi^0\pi^0$	seen	963
$4\pi$	seen	925
$\eta\eta$	seen	803
$K\bar{K}$	seen	837
$\gamma\gamma$	seen	972
$\rho\bar{\rho}$	seen	254

<b><math>f_2(2010)</math></b>	$I^G(J^{PC}) = 0^+(2^+ +)$	
Mass $m = 2011^{+60}_{-80}$ MeV		
Full width $\Gamma = 202 \pm 60$ MeV		

<b><math>f_2(2010)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\phi\phi$	seen	†
$K\bar{K}$	seen	876

**$a_4(2040)$**

$I^G(J^{PC}) = 1^-(4^{++})$

Mass  $m = 1995^{+10}_{-8}$  MeV    (S = 1.1)

Full width  $\Gamma = 257^{+25}_{-23}$  MeV    (S = 1.3)

<b><math>a_4(2040)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\bar{K}$	seen	867
$\pi^+\pi^-\pi^0$	seen	973
$\rho\pi$	seen	841
$f_2(1270)\pi$	seen	579
$\omega\pi^-\pi^0$	seen	818
$\omega\rho$	seen	623
$\eta\pi$	seen	917
$\eta'(958)\pi$	seen	760

**$f_4(2050)$**

$I^G(J^{PC}) = 0^+(4^{++})$

Mass  $m = 2018 \pm 11$  MeV    (S = 2.1)

Full width  $\Gamma = 237 \pm 18$  MeV    (S = 1.9)

<b><math>f_4(2050)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\omega\omega$	seen	637
$\pi\pi$	(17.0±1.5) %	1000
$K\bar{K}$	( 6.8 <sup>+3.4</sup> <sub>-1.8</sub> ) × 10 <sup>-3</sup>	880
$\eta\eta$	( 2.1±0.8) × 10 <sup>-3</sup>	848
$4\pi^0$	< 1.2 %	964
$a_2(1320)\pi$	seen	567

**$\phi(2170)$**

$I^G(J^{PC}) = 0^-(1^{--})$

Mass  $m = 2189 \pm 11$  MeV    (S = 1.8)

Full width  $\Gamma = 79 \pm 14$  MeV

<b><math>\phi(2170)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$e^+e^-$	seen	1095
$\phi f_0(980)$	seen	434
$K^+K^-f_0(980) \rightarrow$ $\quad K^+K^-\pi^+\pi^-$	seen	—
$K^+K^-f_0(980) \rightarrow K^+K^-\pi^0\pi^0$	seen	—
$K^{*0}K^\pm\pi^\mp$	not seen	780
$K^*(892)^0\bar{K}^*(892)^0$	not seen	635

**$f_2(2300)$**

$I^G(J^{PC}) = 0^+(2^{++})$

Mass  $m = 2297 \pm 28$  MeV

Full width  $\Gamma = 149 \pm 40$  MeV

<b><math>f_2(2300)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\phi\phi$	seen	529
$K\bar{K}$	seen	1037
$\gamma\gamma$	seen	1149

**$f_2(2340)$**

$I^G(J^{PC}) = 0^+(2^{++})$

Mass  $m = 2345^{+50}_{-40}$  MeV

Full width  $\Gamma = 322^{+70}_{-60}$  MeV

<b><math>f_2(2340)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\phi\phi$	seen	580
$\eta\eta$	seen	1037

## Meson Summary Table

# STRANGE MESONS ( $S = \pm 1$ , $C = B = 0$ )

 $K^+ = u\bar{s}, K^0 = d\bar{s}, \bar{K}^0 = \bar{d}s, K^- = \bar{u}s$ , similarly for  $K^{*}$ 's

 **$K^\pm$** 

$$I(J^P) = \frac{1}{2}(0^-)$$

Mass  $m = 493.677 \pm 0.016$  MeV [ $u$ ] ( $S = 2.8$ )

Mean life  $\tau = (1.2380 \pm 0.0020) \times 10^{-8}$  s ( $S = 1.8$ )

 $c\tau = 3.711$  m

**CPT violation parameters ( $\Delta$  = rate difference/sum)**

$$\Delta(K^\pm \rightarrow \mu^\pm \nu_\mu) = (-0.27 \pm 0.21)\%$$

$$\Delta(K^\pm \rightarrow \pi^\pm \pi^0) = (0.4 \pm 0.6)\% [v]$$

**CP violation parameters ( $\Delta$  = rate difference/sum)**

$$\Delta(K^\pm \rightarrow \pi^\pm e^+ e^-) = (-2.2 \pm 1.6) \times 10^{-2}$$

$$\Delta(K^\pm \rightarrow \pi^\pm \mu^+ \mu^-) = 0.010 \pm 0.023$$

$$\Delta(K^\pm \rightarrow \pi^\pm \pi^0 \gamma) = (0.0 \pm 1.2) \times 10^{-3}$$

$$\Delta(K^\pm \rightarrow \pi^\pm \pi^+ \pi^-) = (0.04 \pm 0.06)\%$$

$$\Delta(K^\pm \rightarrow \pi^\pm \pi^0 \pi^0) = (-0.02 \pm 0.28)\%$$

**T violation parameters**

$$K^+ \rightarrow \pi^0 \mu^+ \nu_\mu \quad P_T = (-1.7 \pm 2.5) \times 10^{-3}$$

$$K^+ \rightarrow \mu^+ \nu_\mu \gamma \quad P_T = (-0.6 \pm 1.9) \times 10^{-2}$$

$$K^+ \rightarrow \pi^0 \mu^+ \nu_\mu \quad \text{Im}(\xi) = -0.006 \pm 0.008$$

**Slope parameter  $g [x]$** 

(See Particle Listings for quadratic coefficients and alternative parametrization related to  $\pi\pi$  scattering)

$$K^\pm \rightarrow \pi^\pm \pi^+ \pi^- \quad g = -0.21134 \pm 0.00017$$

$$(g_+ - g_-) / (g_+ + g_-) = (-1.5 \pm 2.2) \times 10^{-4}$$

$$K^\pm \rightarrow \pi^\pm \pi^0 \pi^0 \quad g = 0.626 \pm 0.007$$

$$(g_+ - g_-) / (g_+ + g_-) = (1.8 \pm 1.8) \times 10^{-4}$$

 **$K^\pm$  decay form factors [ $a, v$ ]**

Assuming  $\mu$ -e universality

$$\lambda_+(K_{\mu 3}^+) = \lambda_+(K_{e 3}^+) = (2.97 \pm 0.05) \times 10^{-2}$$

$$\lambda_0(K_{\mu 3}^+) = (1.95 \pm 0.12) \times 10^{-2}$$

Not assuming  $\mu$ -e universality

$$\lambda_+(K_{e 3}^+) = (2.98 \pm 0.05) \times 10^{-2}$$

$$\lambda_+(K_{\mu 3}^+) = (2.96 \pm 0.17) \times 10^{-2}$$

$$\lambda_0(K_{\mu 3}^+) = (1.96 \pm 0.13) \times 10^{-2}$$

 $K_{e 3}$  form factor quadratic fit

$$\lambda'_+(K_{e 3}^\pm) \text{ linear coeff.} = (2.49 \pm 0.17) \times 10^{-2}$$

$$\lambda''_+(K_{e 3}^\pm) \text{ quadratic coeff.} = (0.19 \pm 0.09) \times 10^{-2}$$

$$K_{e 3}^+ \quad |f_S/f_+| = (-0.3 \pm 0.8) \times 10^{-2}$$

$$K_{e 3}^+ \quad |f_T/f_+| = (-1.2 \pm 2.3) \times 10^{-2}$$

$$K_{\mu 3}^+ \quad |f_S/f_+| = (0.2 \pm 0.6) \times 10^{-2}$$

$$K_{\mu 3}^+ \quad |f_T/f_+| = (-0.1 \pm 0.7) \times 10^{-2}$$

$$K^+ \rightarrow e^+ \nu_e \gamma \quad |F_A + F_V| = 0.133 \pm 0.008 \quad (S = 1.3)$$

$$K^+ \rightarrow \mu^+ \nu_\mu \gamma \quad |F_A + F_V| = 0.165 \pm 0.013$$

$$K^+ \rightarrow e^+ \nu_e \gamma \quad |F_A - F_V| < 0.49, \text{ CL} = 90\%$$

$$K^+ \rightarrow \mu^+ \nu_\mu \gamma \quad |F_A - F_V| = -0.24 \text{ to } 0.04, \text{ CL} = 90\%$$

**Charge radius**

$$\langle r \rangle = 0.560 \pm 0.031 \text{ fm}$$

**Forward-backward asymmetry**

$$A_{FB}(K_{\pi\mu\mu}^\pm) = \frac{\Gamma(\cos(\theta_{K\mu}) > 0) - \Gamma(\cos(\theta_{K\mu}) < 0)}{\Gamma(\cos(\theta_{K\mu}) > 0) + \Gamma(\cos(\theta_{K\mu}) < 0)} < 2.3 \times 10^{-2}, \text{ CL} = 90\%$$

 $K^-$  modes are charge conjugates of the modes below.

<b><math>K^+</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level (MeV/c)	$p$
<b>Leptonic and semileptonic modes</b>			
$e^+ \nu_e$	( $1.582 \pm 0.007$ ) $\times 10^{-5}$		247
$\mu^+ \nu_\mu$	( $63.56 \pm 0.11$ ) %	$S=1.2$	236
$\pi^0 e^+ \nu_e$	( $5.07 \pm 0.04$ ) %	$S=2.1$	228
Called $K_{e 3}^+$ .			
$\pi^0 \mu^+ \nu_\mu$	( $3.352 \pm 0.033$ ) %	$S=1.9$	215
Called $K_{\mu 3}^+$ .			
$\pi^0 \pi^0 e^+ \nu_e$	( $2.55 \pm 0.04$ ) $\times 10^{-5}$	$S=1.1$	206
$\pi^+ \pi^- e^+ \nu_e$	( $4.247 \pm 0.024$ ) $\times 10^{-5}$		203
$\pi^+ \pi^- \mu^+ \nu_\mu$	( $1.4 \pm 0.9$ ) $\times 10^{-5}$		151
$\pi^0 \pi^0 \pi^0 e^+ \nu_e$	$< 3.5$ $\times 10^{-6}$	CL=90%	135
<b>Hadronic modes</b>			
$\pi^+ \pi^0$	( $20.67 \pm 0.08$ ) %	$S=1.2$	205
$\pi^+ \pi^0 \pi^0$	( $1.760 \pm 0.023$ ) %	$S=1.1$	133
$\pi^+ \pi^+ \pi^-$	( $5.583 \pm 0.024$ ) %		125
<b>Leptonic and semileptonic modes with photons</b>			
$\mu^+ \nu_\mu \gamma$	[z,aa] ( $6.2 \pm 0.8$ ) $\times 10^{-3}$		236
$\mu^+ \nu_\mu \gamma$ (SD $^+$ )	[a,bb] ( $1.33 \pm 0.22$ ) $\times 10^{-5}$		—
$\mu^+ \nu_\mu \gamma$ (SD $^+$ INT)	[a,bb] $< 2.7$ $\times 10^{-5}$	CL=90%	—
$\mu^+ \nu_\mu \gamma$ (SD $^-$ + SD $^-$ INT)	[a,bb] $< 2.6$ $\times 10^{-4}$	CL=90%	—
$e^+ \nu_e \gamma$	( $9.4 \pm 0.4$ ) $\times 10^{-6}$		247
$\pi^0 e^+ \nu_e \gamma$	[z,aa] ( $2.56 \pm 0.16$ ) $\times 10^{-4}$		228
$\pi^0 e^+ \nu_e \gamma$ (SD)	[a,bb] $< 5.3$ $\times 10^{-5}$	CL=90%	228
$\pi^0 \mu^+ \nu_\mu \gamma$	[z,aa] ( $1.25 \pm 0.25$ ) $\times 10^{-5}$		215
$\pi^0 \pi^0 e^+ \nu_e \gamma$	$< 5$ $\times 10^{-6}$	CL=90%	206
<b>Hadronic modes with photons or <math>\ell\ell</math> pairs</b>			
$\pi^+ \pi^0 \gamma$ (INT)	( $-4.2 \pm 0.9$ ) $\times 10^{-6}$		—
$\pi^+ \pi^0 \gamma$ (DE)	[z,cc] ( $6.0 \pm 0.4$ ) $\times 10^{-6}$		205
$\pi^+ \pi^0 \pi^0 \gamma$	[z,aa] ( $7.6 \pm 6.0$ ) $\times 10^{-6}$		133
$\pi^+ \pi^+ \pi^- \gamma$	[z,aa] ( $1.04 \pm 0.31$ ) $\times 10^{-4}$		125
$\pi^+ \gamma \gamma$	[z] ( $1.01 \pm 0.06$ ) $\times 10^{-6}$		227
$\pi^+ 3\gamma$	[z] $< 1.0$ $\times 10^{-4}$	CL=90%	227
$\pi^+ e^+ e^- \gamma$	( $1.19 \pm 0.13$ ) $\times 10^{-8}$		227
<b>Leptonic modes with <math>\ell\ell</math> pairs</b>			
$e^+ \nu_e \nu_e \bar{\nu}_e$	$< 6$ $\times 10^{-5}$	CL=90%	247
$\mu^+ \nu_\mu \nu_\mu \bar{\nu}_\mu$	$< 6.0$ $\times 10^{-6}$	CL=90%	236
$e^+ \nu_e e^+ e^-$	( $2.48 \pm 0.20$ ) $\times 10^{-8}$		247
$\mu^+ \nu_\mu e^+ e^-$	( $7.06 \pm 0.31$ ) $\times 10^{-8}$		236
$e^+ \nu_e \mu^+ \mu^-$	( $1.7 \pm 0.5$ ) $\times 10^{-8}$		223
$\mu^+ \nu_\mu \mu^+ \mu^-$	$< 4.1$ $\times 10^{-7}$	CL=90%	185
<b>Lepton family number (LF), Lepton number (L), <math>\Delta S = \Delta Q</math> (SQ) violating modes, or <math>\Delta S = 1</math> weak neutral current (S1) modes</b>			
$\pi^+ \pi^+ e^- \bar{\nu}_e$	SQ $< 1.3$ $\times 10^{-8}$	CL=90%	203
$\pi^+ \pi^+ \mu^- \bar{\nu}_\mu$	SQ $< 3.0$ $\times 10^{-6}$	CL=95%	151
$\pi^+ e^+ e^-$	S1 ( $3.00 \pm 0.09$ ) $\times 10^{-7}$		227
$\pi^+ \mu^+ \mu^-$	S1 ( $9.4 \pm 0.6$ ) $\times 10^{-8}$	$S=2.6$	172
$\pi^+ \nu \bar{\nu}$	S1 ( $1.7 \pm 1.1$ ) $\times 10^{-10}$		227
$\pi^+ \pi^0 \nu \bar{\nu}$	S1 $< 4.3$ $\times 10^{-5}$	CL=90%	205
$\mu^- \nu e^+ e^+$	LF $< 2.1$ $\times 10^{-8}$	CL=90%	236
$\mu^+ \nu_e$	LF [d] $< 4$ $\times 10^{-3}$	CL=90%	236
$\pi^+ \mu^+ e^-$	LF $< 1.3$ $\times 10^{-11}$	CL=90%	214
$\pi^+ \mu^- e^+$	LF $< 5.2$ $\times 10^{-10}$	CL=90%	214
$\pi^- \mu^+ e^+$	L $< 5.0$ $\times 10^{-10}$	CL=90%	214
$\pi^- e^+ e^+$	L $< 6.4$ $\times 10^{-10}$	CL=90%	227
$\pi^- \mu^+ \mu^+$	L [d] $< 1.1$ $\times 10^{-9}$	CL=90%	172
$\mu^+ \bar{\nu}_e$	L [d] $< 3.3$ $\times 10^{-3}$	CL=90%	236
$\pi^0 e^+ \bar{\nu}_e$	L $< 3$ $\times 10^{-3}$	CL=90%	228
$\pi^+ \gamma$	[dd] $< 2.3$ $\times 10^{-9}$	CL=90%	227

 **$K^0$** 

$$I(J^P) = \frac{1}{2}(0^-)$$

50%  $K_S$ , 50%  $K_L$ 

Mass  $m = 497.611 \pm 0.013$  MeV ( $S = 1.2$ )

 $m_{K^0} - m_{K^\pm} = 3.934 \pm 0.020$  MeV ( $S = 1.6$ )

**Mean square charge radius**

$$\langle r^2 \rangle = -0.077 \pm 0.010 \text{ fm}^2$$

# Meson Summary Table

## T-violation parameters in $K^0\text{--}\bar{K}^0$ mixing [v]

Asymmetry  $A_T$  in  $K^0\text{--}\bar{K}^0$  mixing =  $(6.6 \pm 1.6) \times 10^{-3}$

## CP-violation parameters

$\text{Re}(\epsilon) = (1.596 \pm 0.013) \times 10^{-3}$

## CPT-violation parameters [v]

$\text{Re } \delta = (2.5 \pm 2.3) \times 10^{-4}$

$\text{Im } \delta = (-1.5 \pm 1.6) \times 10^{-5}$

$\text{Re}(y)$ ,  $K_{e3}$  parameter =  $(0.4 \pm 2.5) \times 10^{-3}$

$\text{Re}(x_-)$ ,  $K_{e3}$  parameter =  $(-2.9 \pm 2.0) \times 10^{-3}$

$|m_{K^0} - m_{\bar{K}^0}| / m_{\text{average}} < 6 \times 10^{-19}$ , CL = 90% [ee]

$(\Gamma_{K^0} - \Gamma_{\bar{K}^0}) / m_{\text{average}} = (8 \pm 8) \times 10^{-18}$

## Tests of $\Delta S = \Delta Q$

$\text{Re}(x_+)$ ,  $K_{e3}$  parameter =  $(-0.9 \pm 3.0) \times 10^{-3}$

**$K_S^0$**

$$I(J^P) = \frac{1}{2}(0^-)$$

Mean life  $\tau = (0.8954 \pm 0.0004) \times 10^{-10}$  s (S = 1.1) Assuming CPT

Mean life  $\tau = (0.89564 \pm 0.00033) \times 10^{-10}$  s Not assuming CPT

$c\tau = 2.6844$  cm Assuming CPT

## CP-violation parameters [ff]

$\text{Im}(\eta_{+-0}) = -0.002 \pm 0.009$

$\text{Im}(\eta_{000}) = -0.001 \pm 0.016$

$|\eta_{000}| = |A(K_S^0 \rightarrow 3\pi^0)/A(K_L^0 \rightarrow 3\pi^0)| < 0.0088$ , CL = 90%

CP asymmetry  $A$  in  $\pi^+\pi^-\pi^+\pi^- = (-0.4 \pm 0.8)\%$

$K_S^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
<b>Hadronic modes</b>			
$\pi^0\pi^0$	$(30.69 \pm 0.05)\%$		209
$\pi^+\pi^-$	$(69.20 \pm 0.05)\%$		206
$\pi^+\pi^-\pi^0$	$(3.5 \pm 1.1 \pm 0.9) \times 10^{-7}$		133
<b>Modes with photons or <math>\ell\bar{\ell}</math> pairs</b>			
$\pi^+\pi^-\gamma$	[aa,gg] $(1.79 \pm 0.05) \times 10^{-3}$		206
$\pi^+\pi^-\pi^0 e^+e^-$	$(4.79 \pm 0.15) \times 10^{-5}$		206
$\pi^0\gamma\gamma$	[gg] $(4.9 \pm 1.8) \times 10^{-8}$		230
$\gamma\gamma$	$(2.63 \pm 0.17) \times 10^{-6}$	S=3.0	249
<b>Semileptonic modes</b>			
$\pi^\pm e^\mp \nu_e$	[hh] $(7.04 \pm 0.08) \times 10^{-4}$		229
<b>CP violating (CP) and <math>\Delta S = 1</math> weak neutral current (S1) modes</b>			
$3\pi^0$	CP $< 2.6 \times 10^{-8}$	CL=90%	139
$\mu^+\mu^-$	S1 $< 9 \times 10^{-9}$	CL=90%	225
$e^+e^-$	S1 $< 9 \times 10^{-9}$	CL=90%	249
$\pi^0 e^+ e^-$	S1 [gg] $(3.0 \pm 1.5 \pm 1.2) \times 10^{-9}$		230
$\pi^0 \mu^+ \mu^-$	S1 $(2.9 \pm 1.5 \pm 1.2) \times 10^{-9}$		177

**$K_L^0$**

$$I(J^P) = \frac{1}{2}(0^-)$$

$m_{K_L} - m_{K_S}$

$= (0.5293 \pm 0.0009) \times 10^{10} \hbar \text{ s}^{-1}$  (S = 1.3) Assuming CPT

$= (3.484 \pm 0.006) \times 10^{-12} \text{ MeV}$  Assuming CPT

$= (0.5289 \pm 0.0010) \times 10^{10} \hbar \text{ s}^{-1}$  Not assuming CPT

Mean life  $\tau = (5.116 \pm 0.021) \times 10^{-8} \text{ s}$  (S = 1.1)

$c\tau = 15.34 \text{ m}$

## Slope parameters [x]

(See Particle Listings for other linear and quadratic coefficients)

$K_L^0 \rightarrow \pi^+\pi^-\pi^0$ :  $g = 0.678 \pm 0.008$  (S = 1.5)

$K_L^0 \rightarrow \pi^+\pi^-\pi^0$ :  $h = 0.076 \pm 0.006$

$K_L^0 \rightarrow \pi^+\pi^-\pi^0$ :  $k = 0.0099 \pm 0.0015$

$K_L^0 \rightarrow \pi^0\pi^0\pi^0$ :  $h = (0.6 \pm 1.2) \times 10^{-3}$

## $K_L$ decay form factors [v]

Linear parametrization assuming  $\mu$ -e universality

$\lambda_+(K_{\mu 3}^0) = \lambda_+(K_{e 3}^0) = (2.82 \pm 0.04) \times 10^{-2}$  (S = 1.1)

$\lambda_0(K_{\mu 3}^0) = (1.38 \pm 0.18) \times 10^{-2}$  (S = 2.2)

Quadratic parametrization assuming  $\mu$ -e universality

$\lambda'_+(K_{\mu 3}^0) = \lambda'_+(K_{e 3}^0) = (2.40 \pm 0.12) \times 10^{-2}$  (S = 1.2)

$\lambda''_+(K_{\mu 3}^0) = \lambda''_+(K_{e 3}^0) = (0.20 \pm 0.05) \times 10^{-2}$  (S = 1.2)

$\lambda_0(K_{\mu 3}^0) = (1.16 \pm 0.09) \times 10^{-2}$  (S = 1.2)

Pole parametrization assuming  $\mu$ -e universality

$M_V^\mu(K_{\mu 3}^0) = M_V^e(K_{e 3}^0) = 878 \pm 6 \text{ MeV}$  (S = 1.1)

$M_S^\mu(K_{\mu 3}^0) = 1252 \pm 90 \text{ MeV}$  (S = 2.6)

Dispersive parametrization assuming  $\mu$ -e universality

$\Lambda_+ = (0.251 \pm 0.006) \times 10^{-1}$  (S = 1.5)

$\ln(C) = (1.75 \pm 0.18) \times 10^{-1}$  (S = 2.0)

$K_{e 3}^0$   $|f_S/f_+| = (1.5 \pm 1.4 \pm 1.6) \times 10^{-2}$

$K_{e 3}^0$   $|f_T/f_+| = (5 \pm 4 \pm 5) \times 10^{-2}$

$K_{\mu 3}^0$   $|f_T/f_+| = (12 \pm 12) \times 10^{-2}$

$K_L \rightarrow \ell^+\ell^-\gamma$ ,  $K_L \rightarrow \ell^+\ell^-\ell'^+\ell'^-$ :  $\alpha_{K^*} = -0.205 \pm 0.022$  (S = 1.8)

$K_L^0 \rightarrow \ell^+\ell^-\gamma$ ,  $K_L^0 \rightarrow \ell^+\ell^-\ell'^+\ell'^-$ :  $\alpha_{DIP} = -1.69 \pm 0.08$  (S = 1.7)

$K_L \rightarrow \pi^+\pi^-\pi^+\pi^-$ :  $a_1/a_2 = -0.737 \pm 0.014 \text{ GeV}^2$

$K_L \rightarrow \pi^0 2\gamma$ :  $a_V = -0.43 \pm 0.06$  (S = 1.5)

## CP-violation parameters [ff]

$A_L = (0.332 \pm 0.006)\%$

$|\eta_{00}| = (2.220 \pm 0.011) \times 10^{-3}$  (S = 1.8)

$|\eta_{+-}| = (2.232 \pm 0.011) \times 10^{-3}$  (S = 1.8)

$|\epsilon| = (2.228 \pm 0.011) \times 10^{-3}$  (S = 1.8)

$|\eta_{00}/\eta_{+-}| = 0.9950 \pm 0.0007$  [ii] (S = 1.6)

$\text{Re}(\epsilon'/\epsilon) = (1.66 \pm 0.23) \times 10^{-3}$  [ii] (S = 1.6)

Assuming CPT

$\phi_{+-} = (43.51 \pm 0.05)^\circ$  (S = 1.2)

$\phi_{00} = (43.52 \pm 0.05)^\circ$  (S = 1.3)

$\phi_\epsilon = \phi_{\text{SW}} = (43.52 \pm 0.05)^\circ$  (S = 1.2)

$\text{Im}(\epsilon'/\epsilon) = -(\phi_{00} - \phi_{+-})/3 = (-0.002 \pm 0.005)^\circ$  (S = 1.7)

Not assuming CPT

$\phi_{+-} = (43.4 \pm 0.5)^\circ$  (S = 1.2)

$\phi_{00} = (43.7 \pm 0.6)^\circ$  (S = 1.2)

$\phi_\epsilon = (43.5 \pm 0.5)^\circ$  (S = 1.3)

CP asymmetry  $A$  in  $K_L^0 \rightarrow \pi^+\pi^-\pi^+\pi^- = (13.7 \pm 1.5)\%$

$\beta_{CP}$  from  $K_L^0 \rightarrow e^+e^-\pi^+\pi^- = -0.19 \pm 0.07$

$\gamma_{CP}$  from  $K_L^0 \rightarrow e^+e^-\pi^+\pi^- = 0.01 \pm 0.11$  (S = 1.6)

$j$  for  $K_L^0 \rightarrow \pi^+\pi^-\pi^0 = 0.0012 \pm 0.0008$

$f$  for  $K_L^0 \rightarrow \pi^+\pi^-\pi^0 = 0.004 \pm 0.006$

$|\eta_{+-\gamma}| = (2.35 \pm 0.07) \times 10^{-3}$

$\phi_{+-\gamma} = (44 \pm 4)^\circ$

$|\epsilon'_{+-\gamma}|/\epsilon < 0.3$ , CL = 90%

$|g_{E1}|$  for  $K_L^0 \rightarrow \pi^+\pi^-\gamma < 0.21$ , CL = 90%

## T-violation parameters

$\text{Im}(\xi)$  in  $K_{\mu 3}^0 = -0.007 \pm 0.026$

## CPT invariance tests

$\phi_{00} - \phi_{+-} = (0.34 \pm 0.32)^\circ$

$\text{Re}(\frac{2}{3}\eta_{+-} + \frac{1}{3}\eta_{00}) - \frac{A_\mu}{2} = (-3 \pm 35) \times 10^{-6}$

## $\Delta S = -\Delta Q$ in $K_{e 3}^0$ decay

$\text{Re } x = -0.002 \pm 0.006$

$\text{Im } x = 0.0012 \pm 0.0021$

## Meson Summary Table

$K_L^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level (MeV/c)	$p$
<b>Semileptonic modes</b>			
$\pi^\pm e^\mp \nu_e$ Called $K_{e3}^0$ .	[hh] (40.55 $\pm$ 0.11 ) %	S=1.7	229
$\pi^\pm \mu^\mp \nu_\mu$ Called $K_{\mu 3}^0$ .	[hh] (27.04 $\pm$ 0.07 ) %	S=1.1	216
$(\pi \mu \text{ atom}) \nu$	( 1.05 $\pm$ 0.11 ) $\times 10^{-7}$		188
$\pi^0 \pi^\pm e^\mp \nu$	[hh] ( 5.20 $\pm$ 0.11 ) $\times 10^{-5}$		207
$\pi^\pm e^\mp \nu e^+ e^-$	[hh] ( 1.26 $\pm$ 0.04 ) $\times 10^{-5}$		229
<b>Hadronic modes, including Charge conjugation <math>\times</math> Parity Violating (CPV) modes</b>			
$3\pi^0$	(19.52 $\pm$ 0.12 ) %	S=1.6	139
$\pi^+ \pi^- \pi^0$	(12.54 $\pm$ 0.05 ) %		133
$\pi^+ \pi^-$	CPV [jj] ( 1.967 $\pm$ 0.010 ) $\times 10^{-3}$	S=1.5	206
$\pi^0 \pi^0$	CPV ( 8.64 $\pm$ 0.06 ) $\times 10^{-4}$	S=1.8	209
<b>Semileptonic modes with photons</b>			
$\pi^\pm e^\mp \nu_e \gamma$	[aa, hh, kk] ( 3.79 $\pm$ 0.06 ) $\times 10^{-3}$		229
$\pi^\pm \mu^\mp \nu_\mu \gamma$	( 5.65 $\pm$ 0.23 ) $\times 10^{-4}$		216
<b>Hadronic modes with photons or <math>\ell\bar{\ell}</math> pairs</b>			
$\pi^0 \pi^0 \gamma$	< 2.43 $\times 10^{-7}$	CL=90%	209
$\pi^+ \pi^- \gamma$	[aa, kk] ( 4.15 $\pm$ 0.15 ) $\times 10^{-5}$	S=2.8	206
$\pi^+ \pi^- \gamma$ (DE)	( 2.84 $\pm$ 0.11 ) $\times 10^{-5}$	S=2.0	206
$\pi^0 2\gamma$	[kk] ( 1.273 $\pm$ 0.033 ) $\times 10^{-6}$		230
$\pi^0 \gamma e^+ e^-$	( 1.62 $\pm$ 0.17 ) $\times 10^{-8}$		230
<b>Other modes with photons or <math>\ell\bar{\ell}</math> pairs</b>			
$2\gamma$	( 5.47 $\pm$ 0.04 ) $\times 10^{-4}$	S=1.1	249
$3\gamma$	< 7.4 $\times 10^{-8}$	CL=90%	249
$e^+ e^- \gamma$	( 9.4 $\pm$ 0.4 ) $\times 10^{-6}$	S=2.0	249
$\mu^+ \mu^- \gamma$	( 3.59 $\pm$ 0.11 ) $\times 10^{-7}$	S=1.3	225
$e^+ e^- \gamma \gamma$	[kk] ( 5.95 $\pm$ 0.33 ) $\times 10^{-7}$		249
$\mu^+ \mu^- \gamma \gamma$	[kk] ( 1.0 $\pm^{+0.8}_{-0.6}$ ) $\times 10^{-8}$		225
<b>Charge conjugation <math>\times</math> Parity (CP) or Lepton Family number (LF) violating modes, or <math>\Delta S = 1</math> weak neutral current (S1) modes</b>			
$\mu^+ \mu^-$	S1 ( 6.84 $\pm$ 0.11 ) $\times 10^{-9}$		225
$e^+ e^-$	S1 ( 9 $\pm^{+6}_{-4}$ ) $\times 10^{-12}$		249
$\pi^+ \pi^- e^+ e^-$	S1 [kk] ( 3.11 $\pm$ 0.19 ) $\times 10^{-7}$		206
$\pi^0 \pi^0 e^+ e^-$	S1 < 6.6 $\times 10^{-9}$	CL=90%	209
$\pi^0 \pi^0 \mu^+ \mu^-$	S1 < 9.2 $\times 10^{-11}$	CL=90%	57
$\mu^+ \mu^- e^+ e^-$	S1 ( 2.69 $\pm$ 0.27 ) $\times 10^{-9}$		225
$e^+ e^- e^+ e^-$	S1 ( 3.56 $\pm$ 0.21 ) $\times 10^{-8}$		249
$\pi^0 \mu^+ \mu^-$	CP, S1 [ll] < 3.8 $\times 10^{-10}$	CL=90%	177
$\pi^0 e^+ e^-$	CP, S1 [ll] < 2.8 $\times 10^{-10}$	CL=90%	230
$\pi^0 \nu \bar{\nu}$	CP, S1 [nn] < 2.6 $\times 10^{-8}$	CL=90%	230
$\pi^0 \pi^0 \nu \bar{\nu}$	S1 < 8.1 $\times 10^{-7}$	CL=90%	209
$e^\pm \mu^\mp$	LF [hh] < 4.7 $\times 10^{-12}$	CL=90%	238
$e^\pm e^\pm \mu^\mp \mu^\mp$	LF [hh] < 4.12 $\times 10^{-11}$	CL=90%	225
$\pi^0 \mu^\pm e^\mp$	LF [hh] < 7.6 $\times 10^{-11}$	CL=90%	217
$\pi^0 \pi^0 \mu^\pm e^\mp$	LF < 1.7 $\times 10^{-10}$	CL=90%	159

<b><math>K^*(892)</math></b>		$I(J^P) = \frac{1}{2}(1^-)$
$K^*(892)^\pm$ hadroproduced mass $m = 891.66 \pm 0.26$ MeV		
$K^*(892)^\pm$ in $\tau$ decays mass $m = 895.5 \pm 0.8$ MeV		
$K^*(892)^0$ mass $m = 895.81 \pm 0.19$ MeV (S = 1.4)		
$K^*(892)^\pm$ hadroproduced full width $\Gamma = 50.8 \pm 0.9$ MeV		
$K^*(892)^\pm$ in $\tau$ decays full width $\Gamma = 46.2 \pm 1.3$ MeV		
$K^*(892)^0$ full width $\Gamma = 47.4 \pm 0.6$ MeV (S = 2.2)		
$K^*(892)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level ( $p$ ) (MeV/c)
$K\pi$	$\sim 100$ %	289
$K^0 \gamma$	( 2.46 $\pm$ 0.21 ) $\times 10^{-3}$	307
$K^\pm \gamma$	( 9.9 $\pm$ 0.9 ) $\times 10^{-4}$	309
$K\pi\pi$	< 7 $\times 10^{-4}$	95% 223

<b><math>K_1(1270)</math></b>	$I(J^P) = \frac{1}{2}(1^+)$
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Mass  $m = 1272 \pm 7$  MeV [I]  
Full width  $\Gamma = 90 \pm 20$  MeV [I]

$K_1(1270)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$K\rho$	(42 $\pm 6$ ) %	46
$K_0^*(1430)\pi$	(28 $\pm 4$ ) %	†
$K^*(892)\pi$	(16 $\pm 5$ ) %	302
$K\omega$	(11.0 $\pm$ 2.0 ) %	†
$Kf_0(1370)$	( 3.0 $\pm$ 2.0 ) %	†
$\gamma K^0$	seen	539

<b><math>K_1(1400)</math></b>	$I(J^P) = \frac{1}{2}(1^+)$
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Mass  $m = 1403 \pm 7$  MeV  
Full width  $\Gamma = 174 \pm 13$  MeV (S = 1.6)

$K_1(1400)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$K^*(892)\pi$	(94 $\pm 6$ ) %	402
$K\rho$	( 3.0 $\pm$ 3.0 ) %	293
$Kf_0(1370)$	( 2.0 $\pm$ 2.0 ) %	†
$K\omega$	( 1.0 $\pm$ 1.0 ) %	284
$K_0^*(1430)\pi$	not seen	†
$\gamma K^0$	seen	613

<b><math>K^*(1410)</math></b>	$I(J^P) = \frac{1}{2}(1^-)$
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Mass  $m = 1414 \pm 15$  MeV (S = 1.3)  
Full width  $\Gamma = 232 \pm 21$  MeV (S = 1.1)

$K^*(1410)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$K^*(892)\pi$	> 40 %	95%	410
$K\pi$	( 6.6 $\pm$ 1.3 ) %		612
$K\rho$	< 7 %	95%	305
$\gamma K^0$	seen		619

<b><math>K_0^*(1430)^{[00]}</math></b>	$I(J^P) = \frac{1}{2}(0^+)$
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Mass  $m = 1425 \pm 50$  MeV  
Full width  $\Gamma = 270 \pm 80$  MeV

$K_0^*(1430)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$K\pi$	(93 $\pm 10$ ) %	619
$K\eta$	( 8.6 $\pm^{+2.7}_{-3.4}$ ) %	486
$K\eta'(958)$	seen	†

<b><math>K_2^*(1430)</math></b>	$I(J^P) = \frac{1}{2}(2^+)$
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$K_2^*(1430)^\pm$  mass  $m = 1425.6 \pm 1.5$  MeV (S = 1.1)  
 $K_2^*(1430)^0$  mass  $m = 1432.4 \pm 1.3$  MeV  
 $K_2^*(1430)^\pm$  full width  $\Gamma = 98.5 \pm 2.7$  MeV (S = 1.1)  
 $K_2^*(1430)^0$  full width  $\Gamma = 109 \pm 5$  MeV (S = 1.9)

$K_2^*(1430)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$K\pi$	(49.9 $\pm$ 1.2 ) %		619
$K^*(892)\pi$	(24.7 $\pm$ 1.5 ) %		419
$K^*(892)\pi\pi$	(13.4 $\pm$ 2.2 ) %		372
$K\rho$	( 8.7 $\pm$ 0.8 ) %	S=1.2	318
$K\omega$	( 2.9 $\pm$ 0.8 ) %		311
$K^\pm \gamma$	( 2.4 $\pm$ 0.5 ) $\times 10^{-3}$	S=1.1	627
$K\eta$	( 1.5 $\pm^{+3.4}_{-1.0}$ ) $\times 10^{-3}$	S=1.3	486
$K\omega\pi$	< 7.2 $\times 10^{-4}$	CL=95%	100
$K^0 \gamma$	< 9 $\times 10^{-4}$	CL=90%	626

<b><math>K^*(1680)</math></b>	$I(J^P) = \frac{1}{2}(1^-)$
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Mass  $m = 1717 \pm 27$  MeV (S = 1.4)  
Full width  $\Gamma = 322 \pm 110$  MeV (S = 4.2)

# Meson Summary Table

$K^*(1680)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\pi$	$(38.7 \pm 2.5) \%$	781
$K\rho$	$(31.4^{+5.0}_{-2.1}) \%$	571
$K^*(892)\pi$	$(29.9^{+2.2}_{-5.0}) \%$	618

$K_2(1770)$   $[pp]$

$I(J^P) = \frac{1}{2}(2^-)$

Mass  $m = 1773 \pm 8$  MeV  
Full width  $\Gamma = 186 \pm 14$  MeV

$K_2(1770)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\pi\pi$		794
$K_2^*(1430)\pi$	dominant	288
$K^*(892)\pi$	seen	654
$Kf_2(1270)$	seen	52
$K\phi$	seen	441
$K\omega$	seen	607

$K_3^*(1780)$

$I(J^P) = \frac{1}{2}(3^-)$

Mass  $m = 1776 \pm 7$  MeV (S = 1.1)  
Full width  $\Gamma = 159 \pm 21$  MeV (S = 1.3)

$K_3^*(1780)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$K\rho$	$(31 \pm 9) \%$		613
$K^*(892)\pi$	$(20 \pm 5) \%$		656
$K\pi$	$(18.8 \pm 1.0) \%$		813
$K\eta$	$(30 \pm 13) \%$		719
$K_2^*(1430)\pi$	$< 16 \%$	95%	291

$K_2(1820)$   $[qq]$

$I(J^P) = \frac{1}{2}(2^-)$

Mass  $m = 1816 \pm 13$  MeV  
Full width  $\Gamma = 276 \pm 35$  MeV

$K_2(1820)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K_2^*(1430)\pi$	seen	327
$K^*(892)\pi$	seen	681
$Kf_2(1270)$	seen	185
$K\omega$	seen	638

$K_4^*(2045)$

$I(J^P) = \frac{1}{2}(4^+)$

Mass  $m = 2045 \pm 9$  MeV (S = 1.1)  
Full width  $\Gamma = 198 \pm 30$  MeV

$K_4^*(2045)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$K\pi$	$(9.9 \pm 1.2) \%$	958
$K^*(892)\pi\pi$	$(9 \pm 5) \%$	802
$K^*(892)\pi\pi\pi$	$(7 \pm 5) \%$	768
$\rho K\pi$	$(5.7 \pm 3.2) \%$	741
$\omega K\pi$	$(5.0 \pm 3.0) \%$	738
$\phi K\pi$	$(2.8 \pm 1.4) \%$	594
$\phi K^*(892)$	$(1.4 \pm 0.7) \%$	363

## CHARMED MESONS (C = ±1)

$D^+ = c\bar{d}, D^0 = c\bar{u}, \bar{D}^0 = \bar{c}u, D^- = \bar{c}d,$  similarly for  $D^{*}$ 's

$D^\pm$

$I(J^P) = \frac{1}{2}(0^-)$

Mass  $m = 1869.58 \pm 0.09$  MeV  
Mean life  $\tau = (1040 \pm 7) \times 10^{-15}$  s  
 $c\tau = 311.8 \mu\text{m}$

### c-quark decays

$\Gamma(c \rightarrow \ell^+ \text{ anything})/\Gamma(c \rightarrow \text{ anything}) = 0.096 \pm 0.004$   $[rr]$   
 $\Gamma(c \rightarrow D^*(2010)^+ \text{ anything})/\Gamma(c \rightarrow \text{ anything}) = 0.255 \pm 0.017$

### CP-violation decay-rate asymmetries

$A_{CP}(\mu^\pm \nu) = (8 \pm 8) \%$   
 $A_{CP}(K^0 e^\pm \nu) = (-0.6 \pm 1.6) \%$   
 $A_{CP}(K_S^0 \pi^\pm) = (-0.41 \pm 0.09) \%$   
 $A_{CP}(K^\mp 2\pi^\pm) = (-0.18 \pm 0.16) \%$   
 $A_{CP}(K^\mp \pi^\pm \pi^\pm \pi^0) = (-0.3 \pm 0.7) \%$   
 $A_{CP}(K_S^0 \pi^\pm \pi^0) = (-0.1 \pm 0.7) \%$   
 $A_{CP}(K_S^0 \pi^\pm \pi^+ \pi^-) = (0.0 \pm 1.2) \%$   
 $A_{CP}(\pi^\pm \pi^0) = (2.9 \pm 2.9) \%$   
 $A_{CP}(\pi^\pm \eta) = (1.0 \pm 1.5) \%$  (S = 1.4)  
 $A_{CP}(\pi^\pm \eta'(958)) = (-0.5 \pm 1.2) \%$  (S = 1.1)  
 $A_{CP}(\bar{K}^0/K^0 K^\pm) = (0.11 \pm 0.17) \%$   
 $A_{CP}(K_S^0 K^\pm) = (-0.11 \pm 0.25) \%$   
 $A_{CP}(K^+ K^- \pi^\pm) = (0.37 \pm 0.29) \%$   
 $A_{CP}(K^\pm K^{*0}) = (-0.3 \pm 0.4) \%$   
 $A_{CP}(\phi \pi^\pm) = (0.09 \pm 0.19) \%$  (S = 1.2)  
 $A_{CP}(K^\pm K_0^*(1430)^0) = (8^{+7}_{-6}) \%$   
 $A_{CP}(K^\pm K_2^*(1430)^0) = (43^{+20}_{-26}) \%$   
 $A_{CP}(K^\pm K_0^*(800)) = (-12^{+18}_{-13}) \%$   
 $A_{CP}(a_0(1450)^0 \pi^\pm) = (-19^{+14}_{-16}) \%$   
 $A_{CP}(\phi(1680) \pi^\pm) = (-9 \pm 26) \%$   
 $A_{CP}(\pi^+ \pi^- \pi^\pm) = (-2 \pm 4) \%$   
 $A_{CP}(K_S^0 K^\pm \pi^+ \pi^-) = (-4 \pm 7) \%$   
 $A_{CP}(K^\pm \pi^0) = (-4 \pm 11) \%$

### $\chi^2$ tests of CP-violation (CPV)

Local CPV in  $D^\pm \rightarrow \pi^+ \pi^- \pi^\pm = 78.1 \%$   
Local CPV in  $D^\pm \rightarrow K^+ K^- \pi^\pm = 31 \%$

### CP violating asymmetries of P-odd (T-odd) moments

$A_T(K_S^0 K^\pm \pi^+ \pi^-) = (-12 \pm 11) \times 10^{-3}$   $[ss]$

### D+ form factors

$f_+(0)|V_{cs}|$  in  $\bar{K}^0 \ell^+ \nu_\ell = 0.725 \pm 0.015$  (S = 1.7)  
 $r_1 \equiv a_1/a_0$  in  $\bar{K}^0 \ell^+ \nu_\ell = -1.8 \pm 0.4$   
 $r_2 \equiv a_2/a_0$  in  $\bar{K}^0 \ell^+ \nu_\ell = -3 \pm 12$  (S = 1.5)  
 $f_+(0)|V_{cd}|$  in  $\pi^0 \ell^+ \nu_\ell = 0.146 \pm 0.007$   
 $r_1 \equiv a_1/a_0$  in  $\pi^0 \ell^+ \nu_\ell = -1.4 \pm 0.9$   
 $r_2 \equiv a_2/a_0$  in  $\pi^0 \ell^+ \nu_\ell = -4 \pm 5$   
 $f_+(0)|V_{cd}|$  in  $D^+ \rightarrow \eta e^+ \nu_e = 0.086 \pm 0.006$   
 $r_1 \equiv a_1/a_0$  in  $D^+ \rightarrow \eta e^+ \nu_e = -1.8 \pm 2.2$   
 $r_v \equiv V(0)/A_1(0)$  in  $D^+ \rightarrow \omega e^+ \nu_e = 1.24 \pm 0.11$   
 $r_2 \equiv A_2(0)/A_1(0)$  in  $D^+ \rightarrow \omega e^+ \nu_e = 1.06 \pm 0.16$   
 $r_v \equiv V(0)/A_1(0)$  in  $D^+, D^0 \rightarrow \rho e^+ \nu_e = 1.48 \pm 0.16$   
 $r_2 \equiv A_2(0)/A_1(0)$  in  $D^+, D^0 \rightarrow \rho e^+ \nu_e = 0.83 \pm 0.12$   
 $r_v \equiv V(0)/A_1(0)$  in  $\bar{K}^*(892)^0 \ell^+ \nu_\ell = 1.51 \pm 0.07$  (S = 2.2)  
 $r_2 \equiv A_2(0)/A_1(0)$  in  $\bar{K}^*(892)^0 \ell^+ \nu_\ell = 0.807 \pm 0.025$   
 $r_3 \equiv A_3(0)/A_1(0)$  in  $\bar{K}^*(892)^0 \ell^+ \nu_\ell = 0.0 \pm 0.4$   
 $\Gamma_L/\Gamma_T$  in  $\bar{K}^*(892)^0 \ell^+ \nu_\ell = 1.13 \pm 0.08$   
 $\Gamma_+/ \Gamma_-$  in  $\bar{K}^*(892)^0 \ell^+ \nu_\ell = 0.22 \pm 0.06$  (S = 1.6)

## Meson Summary Table

Most decay modes (other than the semileptonic modes) that involve a neutral  $K$  meson are now given as  $K_S^0$  modes, not as  $\bar{K}^0$  modes. Nearly always it is a  $K_S^0$  that is measured, and interference between Cabibbo-allowed and doubly Cabibbo-suppressed modes can invalidate the assumption that  $2\Gamma(K_S^0) = \Gamma(\bar{K}^0)$ .

<b>D<sup>+</sup> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
<b>Inclusive modes</b>			
$e^+$ semileptonic	$(16.07 \pm 0.30) \%$	—	—
$\mu^+$ anything	$(17.6 \pm 3.2) \%$	—	—
$K^-$ anything	$(25.7 \pm 1.4) \%$	—	—
$\bar{K}^0$ anything + $K^0$ anything	$(61 \pm 5) \%$	—	—
$K^+$ anything	$(5.9 \pm 0.8) \%$	—	—
$K^*(892)^-$ anything	$(6 \pm 5) \%$	—	—
$\bar{K}^*(892)^0$ anything	$(23 \pm 5) \%$	—	—
$K^*(892)^0$ anything	$< 6.6 \%$	CL=90%	—
$\eta$ anything	$(6.3 \pm 0.7) \%$	—	—
$\eta'$ anything	$(1.04 \pm 0.18) \%$	—	—
$\phi$ anything	$(1.03 \pm 0.12) \%$	—	—
<b>Leptonic and semileptonic modes</b>			
$e^+ \nu_e$	$< 8.8 \times 10^{-6}$	CL=90%	935
$\mu^+ \nu_\mu$	$(3.74 \pm 0.17) \times 10^{-4}$	—	932
$\tau^+ \nu_\tau$	$< 1.2 \times 10^{-3}$	CL=90%	90
$\bar{K}^0 e^+ \nu_e$	$(8.90 \pm 0.15) \%$	—	869
$\bar{K}^0 \mu^+ \nu_\mu$	$(9.3 \pm 0.7) \%$	—	865
$K^- \pi^+ e^+ \nu_e$	$(3.91 \pm 0.11) \%$	—	864
$\bar{K}^*(892)^0 e^+ \nu_e, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(3.68 \pm 0.10) \%$	—	722
$(K^- \pi^+)_{S\text{-wave}} e^+ \nu_e$	$(2.26 \pm 0.11) \times 10^{-3}$	—	—
$\bar{K}^*(1430)^0 e^+ \nu_e, \bar{K}^*(1430)^0 \rightarrow K^- \pi^+$	$< 6 \times 10^{-3}$	CL=90%	—
$\bar{K}_2^*(1430)^0 e^+ \nu_e, \bar{K}_2^*(1430)^0 \rightarrow K^- \pi^+$	$< 5 \times 10^{-4}$	CL=90%	—
$K^- \pi^+ \mu^+ \nu_\mu$ nonresonant	$< 7 \times 10^{-3}$	CL=90%	864
$K^- \pi^+ \mu^+ \nu_\mu$	$(3.9 \pm 0.4) \%$	—	851
$\bar{K}^*(892)^0 \mu^+ \nu_\mu, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(3.52 \pm 0.10) \%$	—	717
$K^- \pi^+ \mu^+ \nu_\mu$ nonresonant	$(2.1 \pm 0.5) \times 10^{-3}$	—	851
$K^- \pi^+ \pi^0 \mu^+ \nu_\mu$	$< 1.6 \times 10^{-3}$	CL=90%	825
$\pi^0 e^+ \nu_e$	$(4.05 \pm 0.18) \times 10^{-3}$	—	930
$\eta e^+ \nu_e$	$(1.14 \pm 0.10) \times 10^{-3}$	—	855
$\rho^0 e^+ \nu_e$	$(2.18 \pm 0.25) \times 10^{-3}$	—	774
$\rho^0 \mu^+ \nu_\mu$	$(2.4 \pm 0.4) \times 10^{-3}$	—	770
$\omega e^+ \nu_e$	$(1.69 \pm 0.11) \times 10^{-3}$	—	771
$\eta'(958) e^+ \nu_e$	$(2.2 \pm 0.5) \times 10^{-4}$	—	689
$\phi e^+ \nu_e$	$< 1.3 \times 10^{-5}$	CL=90%	657
Fractions of some of the following modes with resonances have already appeared above as submodes of particular charged-particle modes.			
$\bar{K}^*(892)^0 e^+ \nu_e$	$(5.52 \pm 0.15) \%$	—	722
$\bar{K}^*(892)^0 \mu^+ \nu_\mu$	$(5.30 \pm 0.15) \%$	—	717
$\bar{K}_0^*(1430)^0 \mu^+ \nu_\mu$	$< 2.5 \times 10^{-4}$	CL=90%	380
$\bar{K}^*(1680)^0 \mu^+ \nu_\mu$	$< 1.6 \times 10^{-3}$	CL=90%	105
<b>Hadronic modes with a <math>\bar{K}</math> or <math>\bar{K}K\bar{K}</math></b>			
$K_S^0 \pi^+$	$(1.53 \pm 0.06) \%$	S=2.8	863
$K_L^0 \pi^+$	$(1.46 \pm 0.05) \%$	—	863
$K^- 2\pi^+$	[tt] $(9.46 \pm 0.24) \%$	S=2.0	846
$(K^- \pi^+)_{S\text{-wave}} \pi^+$	$(7.58 \pm 0.22) \%$	—	846
$\bar{K}_0^*(1430)^0 \pi^+, \bar{K}_0^*(1430)^0 \rightarrow K^- \pi^+$	[uu] $(1.26 \pm 0.07) \%$	—	382
$\bar{K}^*(892)^0 \pi^+, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(1.05 \pm 0.12) \%$	—	714
$\bar{K}^*(1410)^0 \pi^+, \bar{K}^*(1410)^0 \rightarrow K^- \pi^+$	not seen	—	381
$\bar{K}_2^*(1430)^0 \pi^+, \bar{K}_2^*(1430)^0 \rightarrow K^- \pi^+$	[uu] $(2.3 \pm 0.8) \times 10^{-4}$	—	371
$\bar{K}^*(1680)^0 \pi^+, \bar{K}^*(1680)^0 \rightarrow K^- \pi^+$	[uu] $(2.2 \pm 1.1) \times 10^{-4}$	—	58
$K^- (2\pi^+)_{I=2}$	$(1.47 \pm 0.27) \%$	—	—
$K_S^0 \pi^+ \pi^0$	[tt] $(7.24 \pm 0.17) \%$	—	845
$K_S^0 \rho^+$	$(6.04 \pm 0.60) \%$	—	677
$K_S^0 \rho(1450)^+, \rho^+ \rightarrow \pi^+ \pi^0$	$(1.5 \pm 1.2) \times 10^{-3}$	—	—

$\bar{K}^*(892)^0 \pi^+, \bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0$	$(2.59 \pm 0.31) \times 10^{-3}$	—	714
$\bar{K}_0^*(1430)^0 \pi^+, \bar{K}_0^*(1430)^0 \rightarrow K_S^0 \pi^0$	$(2.7 \pm 0.9) \times 10^{-3}$	—	—
$\bar{K}_0^*(1680)^0 \pi^+, \bar{K}_0^*(1680)^0 \rightarrow K_S^0 \pi^0$	$(9 \pm 7) \times 10^{-4}$	—	—
$\bar{K}^0 \pi^+, \bar{K}^0 \rightarrow K_S^0 \pi^0$	$(6 \pm 5) \times 10^{-3}$	—	—
$K_S^0 \pi^+ \pi^0$ nonresonant	$(3 \pm 4) \times 10^{-3}$	—	845
$K_S^0 \pi^+ \pi^0$ nonresonant and $\bar{K}^0 \pi^+$	$(1.35 \pm 0.21) \%$	—	—
$(K_S^0 \pi^0)_{S\text{-wave}} \pi^+$	$(1.25 \pm 0.33) \%$	—	845
$K^- 2\pi^+ \pi^0$	[vv] $(6.14 \pm 0.16) \%$	—	816
$K_S^0 2\pi^+ \pi^-$	[vv] $(3.05 \pm 0.09) \%$	—	814
$K^- 3\pi^+ \pi^-$	[tt] $(5.8 \pm 0.5) \times 10^{-3}$	S=1.1	772
$\bar{K}^*(892)^0 2\pi^+ \pi^-, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(1.2 \pm 0.4) \times 10^{-3}$	—	645
$\bar{K}^*(892)^0 \rho^0 \pi^+, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(2.3 \pm 0.4) \times 10^{-3}$	—	239
$\bar{K}^*(892)^0 a_1(1260)^+$	[xx] $(9.4 \pm 1.9) \times 10^{-3}$	—	†
$K^- \rho^0 2\pi^+$	$(1.74 \pm 0.28) \times 10^{-3}$	—	524
$K^- 3\pi^+ \pi^-$ nonresonant	$(4.1 \pm 3.0) \times 10^{-4}$	—	772
$K^+ 2K_S^0$	$(4.6 \pm 2.1) \times 10^{-3}$	—	545
$K^+ K^- K_S^0 \pi^+$	$(2.3 \pm 0.5) \times 10^{-4}$	—	436
<b>Pionic modes</b>			
$\pi^+ \pi^0$	$(1.24 \pm 0.06) \times 10^{-3}$	—	925
$2\pi^+ \pi^-$	$(3.29 \pm 0.20) \times 10^{-3}$	—	909
$\rho^0 \pi^+$	$(8.4 \pm 1.5) \times 10^{-4}$	—	767
$\pi^+ (\pi^+ \pi^-)_{S\text{-wave}}$	$(1.85 \pm 0.17) \times 10^{-3}$	—	909
$\sigma \pi^+, \sigma \rightarrow \pi^+ \pi^-$	$(1.39 \pm 0.12) \times 10^{-3}$	—	—
$f_0(980) \pi^+, f_0(980) \rightarrow \pi^+ \pi^-$	$(1.58 \pm 0.34) \times 10^{-4}$	—	669
$f_0(1370) \pi^+, f_0(1370) \rightarrow \pi^+ \pi^-$	$(8 \pm 4) \times 10^{-5}$	—	—
$f_2(1270) \pi^+, f_2(1270) \rightarrow \pi^+ \pi^-$	$(5.1 \pm 0.9) \times 10^{-4}$	—	485
$\rho(1450)^0 \pi^+, \rho(1450)^0 \rightarrow \pi^+ \pi^-$	$< 8 \times 10^{-5}$	CL=95%	338
$f_0(1500) \pi^+, f_0(1500) \rightarrow \pi^+ \pi^-$	$(1.1 \pm 0.4) \times 10^{-4}$	—	—
$f_0(1710) \pi^+, f_0(1710) \rightarrow \pi^+ \pi^-$	$< 5 \times 10^{-5}$	CL=95%	—
$f_0(1790) \pi^+, f_0(1790) \rightarrow \pi^+ \pi^-$	$< 7 \times 10^{-5}$	CL=95%	—
$(\pi^+ \pi^+)_{S\text{-wave}} \pi^-$	$< 1.2 \times 10^{-4}$	CL=95%	909
$2\pi^+ \pi^-$ nonresonant	$< 1.2 \times 10^{-4}$	CL=95%	909
$\pi^+ 2\pi^0$	$(4.7 \pm 0.4) \times 10^{-3}$	—	910
$2\pi^+ \pi^- \pi^0$	$(1.17 \pm 0.08) \%$	—	883
$\eta \pi^+, \eta \rightarrow \pi^+ \pi^- \pi^0$	$(8.0 \pm 0.5) \times 10^{-4}$	—	848
$\omega \pi^+, \omega \rightarrow \pi^+ \pi^- \pi^0$	$< 3 \times 10^{-4}$	CL=90%	763
$3\pi^+ 2\pi^-$	$(1.67 \pm 0.16) \times 10^{-3}$	—	845

Fractions of some of the following modes with resonances have already appeared above as submodes of particular charged-particle modes.

$\eta \pi^+$	$(3.66 \pm 0.22) \times 10^{-3}$	—	848
$\eta \pi^+ \pi^0$	$(1.38 \pm 0.35) \times 10^{-3}$	—	830
$\omega \pi^+$	$< 3.4 \times 10^{-4}$	CL=90%	764
$\eta'(958) \pi^+$	$(4.84 \pm 0.31) \times 10^{-3}$	—	681
$\eta'(958) \pi^+ \pi^0$	$(1.6 \pm 0.5) \times 10^{-3}$	—	654

Hadronic modes with a  $K\bar{K}$  pair

$K^+ K_S^0$	$(2.95 \pm 0.15) \times 10^{-3}$	S=2.8	793
$K^+ K^- \pi^+$	[tt] $(9.96 \pm 0.26) \times 10^{-3}$	S=1.3	744
$\phi \pi^+, \phi \rightarrow K^+ K^-$	$(2.77 \pm 0.10) \times 10^{-3}$	—	647
$K^+ \bar{K}^*(892)^0, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(2.56 \pm 0.09) \times 10^{-3}$	—	613
$K^+ \bar{K}_0^*(1430)^0, \bar{K}_0^*(1430)^0 \rightarrow K^- \pi^+$	$(1.9 \pm 0.4) \times 10^{-3}$	—	—
$K^+ \bar{K}_2^*(1430)^0, \bar{K}_2^* \rightarrow K^- \pi^+$	$(1.7 \pm 1.3) \times 10^{-4}$	—	—
$K^+ \bar{K}_0^*(800), \bar{K}_0^* \rightarrow K^- \pi^+$	$(7.0 \pm 4.0) \times 10^{-4}$	—	—
$a_0(1450)^0 \pi^+, a_0^0 \rightarrow K^+ K^-$	$(4.6 \pm 7.0) \times 10^{-4}$	—	—

# Meson Summary Table

$\phi(1680)\pi^+, \phi \rightarrow K^+K^-$	$(5.1 \pm 4.0) \times 10^{-5}$	—
$K^+K^-\pi^+$ nonresonant	not seen	744
$K^+K_S^0\pi^+\pi^-$	$(1.71 \pm 0.18) \times 10^{-3}$	678
$K_S^0K^-\pi^+\pi^-$	$(2.34 \pm 0.17) \times 10^{-3}$	678
$K^+K^-\pi^+\pi^-$	$(2.3 \pm 1.2) \times 10^{-4}$	600

A few poorly measured branching fractions:

$\phi\pi^+\pi^0$	$(2.3 \pm 1.0) \%$	619
$\phi\rho^+$	$< 1.5 \%$	CL=90% 260
$K^+K^-\pi^+\pi^0$ non- $\phi$	$(1.5 \pm 0.7) \%$	682
$K^*(892)^+K_S^0$	$(1.7 \pm 0.8) \%$	611

## Doubly Cabibbo-suppressed modes

$K^+\pi^0$	$(1.89 \pm 0.25) \times 10^{-4}$	S=1.2	864
$K^+\eta$	$(1.12 \pm 0.18) \times 10^{-4}$		776
$K^+\eta'(958)$	$(1.83 \pm 0.23) \times 10^{-4}$		571
$K^+\pi^+\pi^-$	$(5.46 \pm 0.25) \times 10^{-4}$		846
$K^+\rho^0$	$(2.1 \pm 0.5) \times 10^{-4}$		679
$K^*(892)^0\pi^+, K^*(892)^0 \rightarrow$	$(2.6 \pm 0.4) \times 10^{-4}$		714
$K^+\pi^-\pi^0, f_0(980) \rightarrow$	$(4.9 \pm 2.9) \times 10^{-5}$		—
$K_2^*(1430)^0\pi^+, K_2^*(1430)^0 \rightarrow$	$(4.4 \pm 3.0) \times 10^{-5}$		—
$K^+\pi^-\pi^0$			
$K^+\pi^+\pi^-$ nonresonant	not seen		846
$2K^+K^-$	$(9.0 \pm 2.1) \times 10^{-5}$		550

## $\Delta C = 1$ weak neutral current (C1) modes, or Lepton Family number (LF) or Lepton number (L) violating modes

$\pi^+e^+e^-$	C1	$< 1.1$	$\times 10^{-6}$	CL=90%	930
$\pi^+\phi, \phi \rightarrow e^+e^-$	[yy]	$(1.7 \pm 1.4) \times 10^{-6}$			—
$\pi^+\mu^+\mu^-$	C1	$< 7.3$	$\times 10^{-8}$	CL=90%	918
$\pi^+\phi, \phi \rightarrow \mu^+\mu^-$	[yy]	$(1.8 \pm 0.8) \times 10^{-6}$			—
$\rho^+\mu^+\mu^-$	C1	$< 5.6$	$\times 10^{-4}$	CL=90%	757
$K^+e^+e^-$	[zz]	$< 1.0$	$\times 10^{-6}$	CL=90%	870
$K^+\mu^+\mu^-$	[zz]	$< 4.3$	$\times 10^{-6}$	CL=90%	856
$\pi^+e^+\mu^-$	LF	$< 2.9$	$\times 10^{-6}$	CL=90%	927
$\pi^+e^-\mu^+$	LF	$< 3.6$	$\times 10^{-6}$	CL=90%	927
$K^+e^+\mu^-$	LF	$< 1.2$	$\times 10^{-6}$	CL=90%	866
$K^+e^-\mu^+$	LF	$< 2.8$	$\times 10^{-6}$	CL=90%	866
$\pi^-2e^+$	L	$< 1.1$	$\times 10^{-6}$	CL=90%	930
$\pi^-2\mu^+$	L	$< 2.2$	$\times 10^{-8}$	CL=90%	918
$\pi^-e^+\mu^+$	L	$< 2.0$	$\times 10^{-6}$	CL=90%	927
$\rho^-2\mu^+$	L	$< 5.6$	$\times 10^{-4}$	CL=90%	757
$K^-2e^+$	L	$< 9$	$\times 10^{-7}$	CL=90%	870
$K^-2\mu^+$	L	$< 1.0$	$\times 10^{-5}$	CL=90%	856
$K^-e^+\mu^+$	L	$< 1.9$	$\times 10^{-6}$	CL=90%	866
$K^*(892)^-2\mu^+$	L	$< 8.5$	$\times 10^{-4}$	CL=90%	703

$$D^0 \quad I(J^P) = \frac{1}{2}(0^-)$$

Mass  $m = 1864.83 \pm 0.05$  MeV  
 $m_{D^\pm} - m_{D^0} = 4.75 \pm 0.08$  MeV  
Mean life  $\tau = (410.1 \pm 1.5) \times 10^{-15}$  s  
 $c\tau = 122.9 \mu\text{m}$

## Mixing and related parameters

$|m_{D_1^0} - m_{D_2^0}| = (0.95^{+0.41}_{-0.44}) \times 10^{10} \hbar \text{ s}^{-1}$   
 $(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma = 2\gamma = (1.29^{+0.14}_{-0.18}) \times 10^{-2}$   
 $|q/p| = 0.92^{+0.12}_{-0.09}$   
 $A_\Gamma = (-0.125 \pm 0.526) \times 10^{-3}$   
 $K^+\pi^-$  relative strong phase:  $\cos \delta = 0.97 \pm 0.11$   
 $K^-\pi^+\pi^0$  coherence factor  $R_{K\pi\pi^0} = 0.82 \pm 0.07$   
 $K^-\pi^+\pi^0$  average relative strong phase  $\delta^{K\pi\pi^0} = (164^{+20}_{-14})^\circ$   
 $K^-\pi^-\pi^+$  coherence factor  $R_{K3\pi} = 0.32^{+0.20}_{-0.28}$   
 $K^-\pi^-\pi^+$  average relative strong phase  $\delta^{K3\pi} = (225^{+21}_{-80})^\circ$   
 $K_S^0K^+\pi^-$  coherence factor  $R_{K_S^0K\pi} = 0.73 \pm 0.08$   
 $K_S^0K^+\pi^-$  average relative strong phase  $\delta^{K_S^0K\pi} = (8 \pm 15)^\circ$   
 $K^*K$  coherence factor  $R_{K^*K} = 1.00 \pm 0.16$   
 $K^*K$  average relative strong phase  $\delta^{K^*K} = (26 \pm 16)^\circ$

## CP-violation decay-rate asymmetries (labeled by the $D^0$ decay)

$A_{CP}(K^+K^-) = (-0.14 \pm 0.12)\%$   
 $A_{CP}(2K_S^0) = (-5 \pm 5)\%$   
 $A_{CP}(\pi^+\pi^-) = (0.01 \pm 0.15)\%$   
 $A_{CP}(2\pi^0) = (0.0 \pm 0.6)\%$   
 $A_{CP}(\pi^+\pi^-\pi^0) = (0.3 \pm 0.4)\%$   
 $A_{CP}(\rho(770)^+\pi^- \rightarrow \pi^+\pi^-\pi^0) = (1.2 \pm 0.9)\% [aaa]$   
 $A_{CP}(\rho(770)^0\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (-3.1 \pm 3.0)\% [aaa]$   
 $A_{CP}(\rho(770)^-\pi^+ \rightarrow \pi^+\pi^-\pi^0) = (-1.0 \pm 1.7)\% [aaa]$   
 $A_{CP}(\rho(1450)^+\pi^- \rightarrow \pi^+\pi^-\pi^0) = (0 \pm 70)\% [aaa]$   
 $A_{CP}(\rho(1450)^0\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (-20 \pm 40)\% [aaa]$   
 $A_{CP}(\rho(1450)^-\pi^+ \rightarrow \pi^+\pi^-\pi^0) = (6 \pm 9)\% [aaa]$   
 $A_{CP}(\rho(1700)^+\pi^- \rightarrow \pi^+\pi^-\pi^0) = (-5 \pm 14)\% [aaa]$   
 $A_{CP}(\rho(1700)^0\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (13 \pm 9)\% [aaa]$   
 $A_{CP}(\rho(1700)^-\pi^+ \rightarrow \pi^+\pi^-\pi^0) = (8 \pm 11)\% [aaa]$   
 $A_{CP}(f_0(980)\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (0 \pm 35)\% [aaa]$   
 $A_{CP}(f_0(1370)\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (25 \pm 18)\% [aaa]$   
 $A_{CP}(f_0(1500)\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (0 \pm 18)\% [aaa]$   
 $A_{CP}(f_0(1710)\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (0 \pm 24)\% [aaa]$   
 $A_{CP}(f_2(1270)\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (-4 \pm 6)\% [aaa]$   
 $A_{CP}(\sigma(400)\pi^0 \rightarrow \pi^+\pi^-\pi^0) = (6 \pm 8)\% [aaa]$   
 $A_{CP}(\text{nonresonant } \pi^+\pi^-\pi^0) = (-13 \pm 23)\% [aaa]$   
 $A_{CP}(K^+K^-\pi^0) = (-1.0 \pm 1.7)\%$   
 $A_{CP}(K^*(892)^+K^- \rightarrow K^+K^-\pi^0) = (-0.9 \pm 1.3)\% [aaa]$   
 $A_{CP}(K^*(1410)^+K^- \rightarrow K^+K^-\pi^0) = (-21 \pm 24)\% [aaa]$   
 $A_{CP}((K^+\pi^0)_{S\text{-wave}}K^- \rightarrow K^+K^-\pi^0) = (7 \pm 15)\% [aaa]$   
 $A_{CP}(\phi(1020)\pi^0 \rightarrow K^+K^-\pi^0) = (1.1 \pm 2.2)\% [aaa]$   
 $A_{CP}(f_0(980)\pi^0 \rightarrow K^+K^-\pi^0) = (-3 \pm 19)\% [aaa]$   
 $A_{CP}(a_0(980)\pi^0 \rightarrow K^+K^-\pi^0) = (-5 \pm 16)\% [aaa]$   
 $A_{CP}(f_2'(1525)\pi^0 \rightarrow K^+K^-\pi^0) = (0 \pm 160)\% [aaa]$   
 $A_{CP}(K^*(892)^-K^+ \rightarrow K^+K^-\pi^0) = (-5 \pm 4)\% [aaa]$   
 $A_{CP}(K^*(1410)^-K^+ \rightarrow K^+K^-\pi^0) = (-17 \pm 29)\% [aaa]$   
 $A_{CP}((K^-\pi^0)_{S\text{-wave}}K^+ \rightarrow K^+K^-\pi^0) = (-10 \pm 40)\% [aaa]$   
 $A_{CP}(K_S^0\pi^0) = (-0.20 \pm 0.17)\%$   
 $A_{CP}(K_S^0\eta) = (0.5 \pm 0.5)\%$   
 $A_{CP}(K_S^0\eta') = (1.0 \pm 0.7)\%$   
 $A_{CP}(K_S^0\phi) = (-3 \pm 9)\%$   
 $A_{CP}(K^-\pi^+) = (0.3 \pm 0.7)\%$   
 $A_{CP}(K^+\pi^-) = (0.0 \pm 1.6)\%$   
 $A_{CP}(D_{CP}(\pm 1) \rightarrow K^\mp\pi^\pm) = (12.7 \pm 1.5)\%$   
 $A_{CP}(K^-\pi^+\pi^0) = (0.1 \pm 0.5)\%$   
 $A_{CP}(K^+\pi^-\pi^0) = (0 \pm 5)\%$   
 $A_{CP}(K_S^0\pi^+\pi^-) = (-0.1 \pm 0.8)\%$   
 $A_{CP}(K^*(892)^-\pi^+ \rightarrow K_S^0\pi^+\pi^-) = (0.4 \pm 0.5)\%$   
 $A_{CP}(K^*(892)^+\pi^- \rightarrow K_S^0\pi^+\pi^-) = (1 \pm 6)\%$   
 $A_{CP}(\overline{K}^0\rho^0 \rightarrow K_S^0\pi^+\pi^-) = (-0.1 \pm 0.5)\%$   
 $A_{CP}(\overline{K}^0\omega \rightarrow K_S^0\pi^+\pi^-) = (-13 \pm 7)\%$   
 $A_{CP}(\overline{K}^0f_0(980) \rightarrow K_S^0\pi^+\pi^-) = (-0.4 \pm 2.7)\%$   
 $A_{CP}(\overline{K}^0f_2(1270) \rightarrow K_S^0\pi^+\pi^-) = (-4 \pm 5)\%$   
 $A_{CP}(\overline{K}^0f_0(1370) \rightarrow K_S^0\pi^+\pi^-) = (-1 \pm 9)\%$   
 $A_{CP}(\overline{K}^0\rho^0(1450) \rightarrow K_S^0\pi^+\pi^-) = (-4 \pm 10)\%$   
 $A_{CP}(\overline{K}^0f_0(600) \rightarrow K_S^0\pi^+\pi^-) = (-3 \pm 5)\%$   
 $A_{CP}(K^*(1410)^-\pi^+ \rightarrow K_S^0\pi^+\pi^-) = (-2 \pm 9)\%$   
 $A_{CP}(K_S^0(1430)^-\pi^+ \rightarrow K_S^0\pi^+\pi^-) = (4 \pm 4)\%$   
 $A_{CP}(K_S^0(1430)^+\pi^- \rightarrow K_S^0\pi^+\pi^-) = (12 \pm 15)\%$   
 $A_{CP}(K_2^*(1430)^-\pi^+ \rightarrow K_S^0\pi^+\pi^-) = (3 \pm 6)\%$   
 $A_{CP}(K_2^*(1430)^+\pi^- \rightarrow K_S^0\pi^+\pi^-) = (-10 \pm 32)\%$   
 $A_{CP}(K^-\pi^+\pi^+\pi^-) = (0.2 \pm 0.5)\%$   
 $A_{CP}(K^+\pi^-\pi^+\pi^-) = (-2 \pm 4)\%$   
 $A_{CP}(K^+K^-\pi^+\pi^-) = (-8 \pm 7)\%$   
 $A_{CP}(K_1^*(1270)^+K^- \rightarrow K^{*0}\pi^+K^-) = (-1 \pm 10)\%$   
 $A_{CP}(K_1^*(1270)^-K^+ \rightarrow \overline{K}^{*0}\pi^-K^+) = (-10 \pm 32)\%$   
 $A_{CP}(K_1^*(1270)^+K^- \rightarrow \rho^0K^+K^-) = (-7 \pm 17)\%$   
 $A_{CP}(K_1^*(1270)^-K^+ \rightarrow \rho^0K^-K^+) = (10 \pm 13)\%$   
 $A_{CP}(K^*(1410)^+K^- \rightarrow K^{*0}\pi^+K^-) = (-20 \pm 17)\%$   
 $A_{CP}(K^*(1410)^-K^+ \rightarrow \overline{K}^{*0}\pi^-K^+) = (-1 \pm 14)\%$   
 $A_{CP}(K^{*0}\overline{K}^{*0} \text{ S-wave}) = (10 \pm 14)\%$   
 $A_{CP}(\phi\rho^0 \text{ S-wave}) = (-3 \pm 5)\%$   
 $A_{CP}(\phi\rho^0 \text{ D-wave}) = (-37 \pm 19)\%$   
 $A_{CP}(\phi(\pi^+\pi^-)_{S\text{-wave}}) = (-9 \pm 10)\%$   
 $A_{CP}((K^-\pi^+)_{P\text{-wave}}(K^+\pi^-)_{S\text{-wave}}) = (3 \pm 11)\%$   
CP-even fraction in  $D^0 \rightarrow \pi^+\pi^-\pi^0$  decays =  $(97.3 \pm 1.7)\%$   
CP-even fraction in  $D^0 \rightarrow K^+K^-\pi^0$  decays =  $(73 \pm 6)\%$   
CP-even fraction in  $D^0 \rightarrow \pi^+\pi^-\pi^+\pi^-$  decays =  $(73.7 \pm 2.8)\%$



## Meson Summary Table

**CP-violation asymmetry difference**

$$\Delta A_{CP} = A_{CP}(K^+ K^-) - A_{CP}(\pi^+ \pi^-) = (-0.32 \pm 0.22)\% \quad (S = 1.9)$$

 **$\chi^2$  tests of CP-violation (CPV)**

$$\begin{aligned} \text{Local CPV in } D^0, \bar{D}^0 \rightarrow \pi^+ \pi^- \pi^0 &= 4.9\% \\ \text{Local CPV in } D^0, \bar{D}^0 \rightarrow \pi^+ \pi^- \pi^+ \pi^- &= 41\% \\ \text{Local CPV in } D^0, \bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^- &= 96\% \\ \text{Local CPV in } D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^0 &= 16.6\% \\ \text{Local CPV in } D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^+ \pi^- &= 9.1\% \end{aligned}$$

**T-violation decay-rate asymmetry**

$$A_T(K^+ K^- \pi^+ \pi^-) = (1.7 \pm 2.7) \times 10^{-3} \text{ [ss]}$$

**CPT-violation decay-rate asymmetry**

$$A_{CPT}(K^\mp \pi^\pm) = 0.008 \pm 0.008$$

**Form factors**

$$\begin{aligned} r_V &\equiv V(0)/A_1(0) \text{ in } D^0 \rightarrow K^*(892)^- \ell^+ \nu_\ell = 1.7 \pm 0.8 \\ r_2 &\equiv A_2(0)/A_1(0) \text{ in } D^0 \rightarrow K^*(892)^- \ell^+ \nu_\ell = 0.9 \pm 0.4 \\ f_+(0) \text{ in } D^0 \rightarrow K^- \ell^+ \nu_\ell &= 0.736 \pm 0.004 \\ f_+(0)|V_{cs}| \text{ in } D^0 \rightarrow K^- \ell^+ \nu_\ell &= 0.719 \pm 0.004 \\ r_1 &\equiv a_1/a_0 \text{ in } D^0 \rightarrow K^- \ell^+ \nu_\ell = -2.40 \pm 0.16 \\ r_2 &\equiv a_2/a_0 \text{ in } D^0 \rightarrow K^- \ell^+ \nu_\ell = 5 \pm 4 \\ f_+(0) \text{ in } D^0 \rightarrow \pi^- \ell^+ \nu_\ell &= 0.637 \pm 0.009 \\ f_+(0)|V_{cd}| \text{ in } D^0 \rightarrow \pi^- \ell^+ \nu_\ell &= 0.1436 \pm 0.0026 \quad (S = 1.5) \\ r_1 &\equiv a_1/a_0 \text{ in } D^0 \rightarrow \pi^- \ell^+ \nu_\ell = -1.97 \pm 0.28 \quad (S = 1.4) \\ r_2 &\equiv a_1/a_0 \text{ in } D^0 \rightarrow \pi^- \ell^+ \nu_\ell = -0.2 \pm 2.2 \quad (S = 1.7) \end{aligned}$$

Most decay modes (other than the semileptonic modes) that involve a neutral  $K$  meson are now given as  $K_S^0$  modes, not as  $\bar{K}^0$  modes. Nearly always it is a  $K_S^0$  that is measured, and interference between Cabibbo-allowed and doubly Cabibbo-suppressed modes can invalidate the assumption that  $2\Gamma(K_S^0) = \Gamma(\bar{K}^0)$ .

<b>D<sup>0</sup> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level (MeV/c)	$p$
<b>Topological modes</b>			
0-prongs	[bbb] (15 $\pm$ 6 ) %	—	—
2-prongs	(70 $\pm$ 6 ) %	—	—
4-prongs	[ccc] (14.5 $\pm$ 0.5 ) %	—	—
6-prongs	[ddd] ( 6.4 $\pm$ 1.3 ) $\times 10^{-4}$	—	—
<b>Inclusive modes</b>			
$e^+$ anything	[eee] ( 6.49 $\pm$ 0.11 ) %	—	—
$\mu^+$ anything	( 6.7 $\pm$ 0.6 ) %	—	—
$K^-$ anything	(54.7 $\pm$ 2.8 ) %	S=1.3	—
$\bar{K}^0$ anything + $K^0$ anything	(47 $\pm$ 4 ) %	—	—
$K^+$ anything	( 3.4 $\pm$ 0.4 ) %	—	—
$K^*(892)^-$ anything	(15 $\pm$ 9 ) %	—	—
$\bar{K}^*(892)^0$ anything	( 9 $\pm$ 4 ) %	—	—
$K^*(892)^+$ anything	< 3.6 %	CL=90%	—
$K^*(892)^0$ anything	( 2.8 $\pm$ 1.3 ) %	—	—
$\eta$ anything	( 9.5 $\pm$ 0.9 ) %	—	—
$\eta'$ anything	( 2.48 $\pm$ 0.27 ) %	—	—
$\phi$ anything	( 1.05 $\pm$ 0.11 ) %	—	—
<b>Semileptonic modes</b>			
$K^- e^+ \nu_e$	( 3.538 $\pm$ 0.033 ) %	S=1.3	867
$K^- \mu^+ \nu_\mu$	( 3.33 $\pm$ 0.13 ) %	—	864
$K^*(892)^- e^+ \nu_e$	( 2.16 $\pm$ 0.16 ) %	—	719
$K^*(892)^- \mu^+ \nu_\mu$	( 1.92 $\pm$ 0.25 ) %	—	714
$K^- \pi^0 e^+ \nu_e$	( 1.6 $\pm$ 1.3 ) %	—	861
$\bar{K}^0 \pi^- e^+ \nu_e$	( 2.7 $\pm$ 0.9 ) %	—	860
$K^- \pi^+ \pi^- e^+ \nu_e$	( 2.8 $\pm$ 1.4 ) $\times 10^{-4}$	—	843
$K_1(1270)^- e^+ \nu_e$	( 7.6 $\pm$ 4.0 ) $\times 10^{-4}$	—	498
$K^- \pi^+ \pi^- \mu^+ \nu_\mu$	< 1.2 $\times 10^{-3}$	CL=90%	821
$(\bar{K}^*(892)\pi)^- \mu^+ \nu_\mu$	< 1.4 $\times 10^{-3}$	CL=90%	692
$\pi^- e^+ \nu_e$	( 2.91 $\pm$ 0.04 ) $\times 10^{-3}$	S=1.1	927
$\pi^- \mu^+ \nu_\mu$	( 2.38 $\pm$ 0.24 ) $\times 10^{-3}$	—	924
$\rho^- e^+ \nu_e$	( 1.77 $\pm$ 0.16 ) $\times 10^{-3}$	—	771

**Hadronic modes with one  $\bar{K}$** 

$K^- \pi^+$	( 3.93 $\pm$ 0.04 ) %	S=1.2	861
$K^+ \pi^-$	( 1.398 $\pm$ 0.027 ) $\times 10^{-4}$	—	861
$K_S^0 \pi^0$	( 1.20 $\pm$ 0.04 ) %	—	860
$K_L^0 \pi^0$	(10.0 $\pm$ 0.7 ) $\times 10^{-3}$	—	860
$K_S^0 \pi^+ \pi^-$	[tt] ( 2.85 $\pm$ 0.20 ) %	S=1.1	842
$K_S^0 \rho^0$	( 6.4 $\pm$ 0.7 ) $\times 10^{-3}$	—	674
$K_S^0 \omega, \omega \rightarrow \pi^+ \pi^-$	( 2.1 $\pm$ 0.6 ) $\times 10^{-4}$	—	670
$K_S^0 (\pi^+ \pi^-)_{S\text{-wave}}$	( 3.4 $\pm$ 0.8 ) $\times 10^{-3}$	—	842
$K_S^0 f_0(980),$ $f_0(980) \rightarrow \pi^+ \pi^-$	( 1.23 $\pm$ 0.40 ) $\times 10^{-3}$	—	549
$K_S^0 f_0(1370),$ $f_0(1370) \rightarrow \pi^+ \pi^-$	( 2.8 $\pm$ 0.9 ) $\times 10^{-3}$	†	—
$K_S^0 f_2(1270),$ $f_2(1270) \rightarrow \pi^+ \pi^-$	( 9 $\pm$ 10 ) $\times 10^{-5}$	—	262
$K^*(892)^- \pi^+,$ $K^*(892)^- \rightarrow K_S^0 \pi^-$	( 1.68 $\pm$ 0.15 ) %	—	711
$K_0^*(1430)^- \pi^+,$ $K_0^*(1430)^- \rightarrow K_S^0 \pi^-$	( 2.73 $\pm$ 0.40 ) $\times 10^{-3}$	—	378
$K_2^*(1430)^- \pi^+,$ $K_2^*(1430)^- \rightarrow K_S^0 \pi^-$	( 3.4 $\pm$ 1.9 ) $\times 10^{-4}$	—	367
$K^*(1680)^- \pi^+,$ $K^*(1680)^- \rightarrow K_S^0 \pi^-$	( 4 $\pm$ 4 ) $\times 10^{-4}$	—	46
$K^*(892)^+ \pi^-,$ $K^*(892)^+ \rightarrow K_S^0 \pi^+$	[fff] ( 1.15 $\pm$ 0.60 ) $\times 10^{-4}$	—	711
$K_0^*(1430)^+ \pi^-,$ $K_0^*(1430)^+ \rightarrow K_S^0 \pi^+$	[fff] < 1.4 $\times 10^{-5}$	CL=95%	—
$K_2^*(1430)^+ \pi^-,$ $K_2^*(1430)^+ \rightarrow K_S^0 \pi^+$	[fff] < 3.4 $\times 10^{-5}$	CL=95%	—
$K_S^0 \pi^+ \pi^-$ nonresonant	( 2.6 $\pm$ 6.0 ) $\times 10^{-4}$	—	842
$K^- \pi^+ \pi^0$	[tt] (14.3 $\pm$ 0.8 ) %	S=3.1	844
$K^- \rho^+$	(11.1 $\pm$ 0.9 ) %	—	675
$K^- \rho(1700)^+,$ $\rho(1700)^+ \rightarrow \pi^+ \pi^0$	( 8.1 $\pm$ 1.8 ) $\times 10^{-3}$	†	—
$K^*(892)^- \pi^+,$ $K^*(892)^- \rightarrow K^- \pi^0$	( 2.28 $\pm$ 0.23 ) %	—	711
$\bar{K}^*(892)^0 \pi^0,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.93 $\pm$ 0.26 ) %	—	711
$K_0^*(1430)^- \pi^+,$ $K_0^*(1430)^- \rightarrow K^- \pi^0$	( 4.7 $\pm$ 2.2 ) $\times 10^{-3}$	—	378
$\bar{K}_0^*(1430)^0 \pi^0,$ $\bar{K}_0^*(1430)^0 \rightarrow K^- \pi^+$	( 5.8 $\pm$ 5.0 ) $\times 10^{-3}$	—	379
$K^*(1680)^- \pi^+,$ $K^*(1680)^- \rightarrow K^- \pi^0$	( 1.9 $\pm$ 0.7 ) $\times 10^{-3}$	—	46
$K^- \pi^+ \pi^0$ nonresonant	( 1.14 $\pm$ 0.50 ) %	—	844
$K_S^0 2\pi^0$	( 9.1 $\pm$ 1.1 ) $\times 10^{-3}$	S=2.2	843
$K_S^0 (2\pi^0)\text{-}S\text{-wave}$	( 2.6 $\pm$ 0.7 ) $\times 10^{-3}$	—	—
$\bar{K}^*(892)^0 \pi^0,$ $\bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0$	( 7.9 $\pm$ 0.7 ) $\times 10^{-3}$	—	711
$\bar{K}^*(1430)^0 \pi^0, \bar{K}^{*0} \rightarrow$ $K_S^0 \pi^0$	( 4 $\pm$ 23 ) $\times 10^{-5}$	—	—
$\bar{K}^*(1680)^0 \pi^0, \bar{K}^{*0} \rightarrow$ $K_S^0 \pi^0$	( 1.0 $\pm$ 0.4 ) $\times 10^{-3}$	—	—
$K_S^0 f_2(1270), f_2 \rightarrow 2\pi^0$	( 2.3 $\pm$ 1.1 ) $\times 10^{-4}$	—	—
$2K_S^0, \text{one } K_S^0 \rightarrow 2\pi^0$	( 3.2 $\pm$ 1.1 ) $\times 10^{-4}$	—	—
$K^- 2\pi^+ \pi^-$	[tt] ( 8.06 $\pm$ 0.23 ) %	S=1.5	813
$K^- \pi^+ \rho^0 \text{ total}$	( 6.73 $\pm$ 0.34 ) %	—	609
$K^- \pi^+ \rho^0 \text{ 3-body}$	( 5.1 $\pm$ 2.3 ) $\times 10^{-3}$	—	609
$\bar{K}^*(892)^0 \rho^0,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.05 $\pm$ 0.23 ) %	—	416
$K^- a_1(1260)^+,$ $a_1(1260)^+ \rightarrow 2\pi^+ \pi^-$	( 3.6 $\pm$ 0.6 ) %	—	327
$\bar{K}^*(892)^0 \pi^+ \pi^- \text{ total},$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.6 $\pm$ 0.4 ) %	—	685
$\bar{K}^*(892)^0 \pi^+ \pi^- \text{ 3-body},$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 9.9 $\pm$ 2.3 ) $\times 10^{-3}$	—	685
$K_1(1270)^- \pi^+,$ $K_1(1270)^- \rightarrow K^- \pi^+ \pi^-$	[ggg] ( 2.9 $\pm$ 0.3 ) $\times 10^{-3}$	—	484
$K^- 2\pi^+ \pi^-$ nonresonant	( 1.88 $\pm$ 0.26 ) %	—	813
$K_S^0 \pi^+ \pi^- \pi^0$	[hhh] ( 5.2 $\pm$ 0.6 ) %	—	813

# Meson Summary Table

$K_S^0 \eta, \eta \rightarrow \pi^+ \pi^- \pi^0$	$(1.02 \pm 0.09) \times 10^{-3}$	772	$\rho(1450)^- \pi^+, \rho(1450)^- \rightarrow \pi^- \pi^0$	$(2.6 \pm 0.4) \times 10^{-4}$	—
$K_S^0 \omega, \omega \rightarrow \pi^+ \pi^- \pi^0$	$(9.9 \pm 0.5) \times 10^{-3}$	670	$\rho(1700)^+ \pi^-, \rho(1700)^+ \rightarrow \pi^+ \pi^0$	$(6.0 \pm 1.5) \times 10^{-4}$	—
$K^- 2\pi^+ \pi^- \pi^0$	$(4.2 \pm 0.4) \%$	771	$\rho(1700)^0 \pi^0, \rho(1700)^0 \rightarrow \pi^+ \pi^-$	$(7.4 \pm 1.8) \times 10^{-4}$	—
$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(1.3 \pm 0.6) \%$	643	$\rho(1700)^- \pi^+, \rho(1700)^- \rightarrow \pi^- \pi^0$	$(4.7 \pm 1.1) \times 10^{-4}$	—
$K^- \pi^+ \omega, \omega \rightarrow \pi^+ \pi^- \pi^0$	$(2.7 \pm 0.5) \%$	605	$f_0(980) \pi^0, f_0(980) \rightarrow \pi^+ \pi^-$	$(3.7 \pm 0.9) \times 10^{-5}$	—
$\bar{K}^*(892)^0 \omega, \bar{K}^*(892)^0 \rightarrow K^- \pi^+, \omega \rightarrow \pi^+ \pi^- \pi^0$	$(6.5 \pm 3.0) \times 10^{-3}$	410	$f_0(500) \pi^0, f_0(500) \rightarrow \pi^+ \pi^-$	$(1.21 \pm 0.22) \times 10^{-4}$	—
$K_S^0 \eta \pi^0$	$(5.5 \pm 1.1) \times 10^{-3}$	721	$f_0(1370) \pi^0, f_0(1370) \rightarrow \pi^+ \pi^-$	$(5.4 \pm 2.1) \times 10^{-5}$	—
$K_S^0 a_0(980), a_0(980) \rightarrow \eta \pi^0$	$(6.6 \pm 2.0) \times 10^{-3}$	—	$f_0(1500) \pi^0, f_0(1500) \rightarrow \pi^+ \pi^-$	$(5.7 \pm 1.6) \times 10^{-5}$	—
$\bar{K}^*(892)^0 \eta, \bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0$	$(1.6 \pm 0.5) \times 10^{-3}$	—	$f_0(1710) \pi^0, f_0(1710) \rightarrow \pi^+ \pi^-$	$(4.6 \pm 1.6) \times 10^{-5}$	—
$K_S^0 2\pi^+ 2\pi^-$	$(2.71 \pm 0.31) \times 10^{-3}$	768	$f_2(1270) \pi^0, f_2(1270) \rightarrow \pi^+ \pi^-$	$(1.94 \pm 0.22) \times 10^{-4}$	—
$K_S^0 \rho^0 \pi^+ \pi^-, \text{no } K^*(892)^-$	$(1.1 \pm 0.7) \times 10^{-3}$	—	$\pi^+ \pi^- \pi^0 \text{ nonresonant}$	$(1.2 \pm 0.4) \times 10^{-4}$	907
$K^*(892)^- 2\pi^+ \pi^-, K^*(892)^- \rightarrow K_S^0 \pi^-, \text{no } \rho^0$	$(5 \pm 8) \times 10^{-4}$	642	$3\pi^0$	$< 3.5 \times 10^{-4}$	CL=90% 908
$K^*(892)^- \rho^0 \pi^+, K^*(892)^- \rightarrow K_S^0 \pi^-$	$(1.6 \pm 0.6) \times 10^{-3}$	230	$2\pi^+ 2\pi^-$	$(7.45 \pm 0.22) \times 10^{-3}$	S=1.2 880
$K_S^0 2\pi^+ 2\pi^- \text{ nonresonant}$	$< 1.2 \times 10^{-3}$	768	$a_1(1260)^+ \pi^-, a_1^+ \rightarrow 2\pi^+ \pi^- \text{ total}$	$(4.47 \pm 0.32) \times 10^{-3}$	—
$K^- 3\pi^+ 2\pi^-$	$(2.2 \pm 0.6) \times 10^{-4}$	713	$a_1(1260)^+ \pi^-, a_1^+ \rightarrow \rho^0 \pi^+ S\text{-wave}$	$(3.23 \pm 0.25) \times 10^{-3}$	—
Fractions of many of the following modes with resonances have already appeared above as submodes of particular charged-particle modes. (Modes for which there are only upper limits and $\bar{K}^*(892)\rho$ submodes only appear below.)					
$K_S^0 \eta$	$(4.85 \pm 0.30) \times 10^{-3}$	772	$a_1(1260)^+ \pi^-, a_1^+ \rightarrow \rho^0 \pi^+ D\text{-wave}$	$(1.9 \pm 0.5) \times 10^{-4}$	—
$K_S^0 \omega$	$(1.11 \pm 0.06) \%$	670	$a_1(1260)^+ \pi^-, a_1^+ \rightarrow \sigma \pi^+ \pi^-$	$(6.2 \pm 0.7) \times 10^{-4}$	—
$K_S^0 \eta'(958)$	$(9.5 \pm 0.5) \times 10^{-3}$	565	$2\rho^0 \text{ total}$	$(1.83 \pm 0.13) \times 10^{-3}$	518
$K^- a_1(1260)^+$	$(7.8 \pm 1.1) \%$	327	$2\rho^0, \text{parallel helicities}$	$(8.2 \pm 3.2) \times 10^{-5}$	—
$K^- a_2(1320)^+$	$< 2 \times 10^{-3}$	CL=90% 198	$2\rho^0, \text{perpendicular helicities}$	$(4.8 \pm 0.6) \times 10^{-4}$	—
$\bar{K}^*(892)^0 \pi^+ \pi^- \text{ total}$	$(2.4 \pm 0.5) \%$	685	$2\rho^0, \text{longitudinal helicities}$	$(1.25 \pm 0.10) \times 10^{-3}$	—
$\bar{K}^*(892)^0 \pi^+ \pi^- 3\text{-body}$	$(1.48 \pm 0.34) \%$	685	$\text{Resonant } (\pi^+ \pi^-) \pi^+ \pi^-$	$(1.49 \pm 0.12) \times 10^{-3}$	—
$\bar{K}^*(892)^0 \rho^0$	$(1.57 \pm 0.35) \%$	417	$3\text{-body total}$	$(6.1 \pm 0.9) \times 10^{-4}$	—
$\bar{K}^*(892)^0 \rho^0 \text{ transverse}$	$(1.7 \pm 0.6) \%$	417	$\sigma \pi^+ \pi^-$	$(1.8 \pm 0.5) \times 10^{-4}$	—
$\bar{K}^*(892)^0 \rho^0 S\text{-wave}$	$(3.0 \pm 0.6) \%$	417	$f_0(980) \pi^+ \pi^-, f_0 \rightarrow \pi^+ \pi^-$	$(3.7 \pm 0.6) \times 10^{-4}$	—
$\bar{K}^*(892)^0 \rho^0 S\text{-wave long.}$	$< 3 \times 10^{-3}$	CL=90% 417	$f_2(1270) \pi^+ \pi^-, f_2 \rightarrow \pi^+ \pi^-$	$(1.01 \pm 0.09) \%$	882
$\bar{K}^*(892)^0 \rho^0 P\text{-wave}$	$< 3 \times 10^{-3}$	CL=90% 417	$\eta \pi^0$	$[iii] (6.9 \pm 0.7) \times 10^{-4}$	846
$\bar{K}^*(892)^0 \rho^0 D\text{-wave}$	$(2.1 \pm 0.6) \%$	417	$\omega \pi^0$	$[iii] < 2.6 \times 10^{-4}$	CL=90% 761
$K_1(1270)^- \pi^+$	$[ggg] (1.6 \pm 0.8) \%$	484	$2\pi^+ 2\pi^- \pi^0$	$(4.2 \pm 0.5) \times 10^{-3}$	844
$K_1(1400)^- \pi^+$	$< 1.2 \%$	CL=90% 386	$\eta \pi^+ \pi^-$	$[iii] (1.09 \pm 0.16) \times 10^{-3}$	827
$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0$	$(1.9 \pm 0.9) \%$	643	$\omega \pi^+ \pi^-$	$[iii] (1.6 \pm 0.5) \times 10^{-3}$	738
$K^- \pi^+ \omega$	$(3.1 \pm 0.6) \%$	605	$3\pi^+ 3\pi^-$	$(4.2 \pm 1.2) \times 10^{-4}$	795
$\bar{K}^*(892)^0 \omega$	$(1.1 \pm 0.5) \%$	410	$\eta'(958) \pi^0$	$(9.1 \pm 1.4) \times 10^{-4}$	678
$K^- \pi^+ \eta'(958)$	$(7.5 \pm 1.9) \times 10^{-3}$	479	$\eta'(958) \pi^+ \pi^-$	$(4.5 \pm 1.7) \times 10^{-4}$	650
$\bar{K}^*(892)^0 \eta'(958)$	$< 1.1 \times 10^{-3}$	CL=90% 119	$2\eta$	$(1.70 \pm 0.20) \times 10^{-3}$	754
<b>Hadronic modes with three <math>K</math>'s</b>					
$K_S^0 K^+ K^-$	$(4.51 \pm 0.34) \times 10^{-3}$	544	$\eta \eta'(958)$	$(1.06 \pm 0.27) \times 10^{-3}$	537
$K_S^0 a_0(980)^0, a_0^0 \rightarrow K^+ K^-$	$(3.0 \pm 0.4) \times 10^{-3}$	—	<b>Hadronic modes with a <math>K\bar{K}</math> pair</b>		
$K^- a_0(980)^+, a_0^+ \rightarrow K^+ K_S^0$	$(6.0 \pm 1.8) \times 10^{-4}$	—	$K^+ K^-$	$(4.01 \pm 0.07) \times 10^{-3}$	S=1.5 791
$K^+ a_0(980)^-, a_0^- \rightarrow K^- K_S^0$	$< 1.1 \times 10^{-4}$	CL=95% —	$2K_S^0$	$(1.8 \pm 0.4) \times 10^{-4}$	S=2.5 789
$K_S^0 f_0(980), f_0 \rightarrow K^+ K^-$	$< 9 \times 10^{-5}$	CL=95% —	$K_S^0 K^- \pi^+$	$(3.6 \pm 0.5) \times 10^{-3}$	S=1.2 739
$K_S^0 \phi, \phi \rightarrow K^+ K^-$	$(2.07 \pm 0.16) \times 10^{-3}$	520	$\bar{K}^*(892)^0 K_S^0, \bar{K}^{*0} \rightarrow K^- \pi^+$	$< 5 \times 10^{-4}$	CL=90% 608
$K_S^0 f_0(1370), f_0 \rightarrow K^+ K^-$	$(1.7 \pm 1.1) \times 10^{-4}$	—	$K_S^0 K^+ \pi^-$	$(2.2 \pm 0.4) \times 10^{-3}$	S=1.3 739
$3K_S^0$	$(9.2 \pm 1.3) \times 10^{-4}$	539	$K^*(892)^0 K_S^0, K^{*0} \rightarrow K^+ \pi^-$	$< 1.8 \times 10^{-4}$	CL=90% 608
$K^+ 2K^- \pi^+$	$(2.21 \pm 0.32) \times 10^{-4}$	434	$K^+ K^- \pi^0$	$(3.38 \pm 0.21) \times 10^{-3}$	743
$K^+ K^- \bar{K}^*(892)^0, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(4.4 \pm 1.7) \times 10^{-5}$	†	$K^*(892)^+ K^-, K^*(892)^+ \rightarrow K^+ \pi^0$	$(1.50 \pm 0.10) \times 10^{-3}$	—
$K^- \pi^+ \phi, \phi \rightarrow K^+ K^-$	$(4.0 \pm 1.7) \times 10^{-5}$	422	$K^*(892)^- K^+, K^*(892)^- \rightarrow K^- \pi^0$	$(5.4 \pm 0.5) \times 10^{-4}$	—
$\phi \bar{K}^*(892)^0, \phi \rightarrow K^+ K^-, \bar{K}^*(892)^0 \rightarrow K^- \pi^+$	$(1.06 \pm 0.20) \times 10^{-4}$	†	$(K^+ \pi^0)_{S\text{-wave}} K^-$	$(2.40 \pm 0.21) \times 10^{-3}$	743
$K^+ 2K^- \pi^+ \text{ nonresonant}$	$(3.3 \pm 1.5) \times 10^{-5}$	434	$(K^- \pi^0)_{S\text{-wave}} K^+$	$(1.3 \pm 0.5) \times 10^{-4}$	743
$2K_S^0 K^\pm \pi^\mp$	$(6.1 \pm 1.3) \times 10^{-4}$	427	$f_0(980) \pi^0, f_0 \rightarrow K^+ K^-$	$(3.5 \pm 0.6) \times 10^{-4}$	—
<b>Pionic modes</b>					
$\pi^+ \pi^-$	$(1.420 \pm 0.025) \times 10^{-3}$	S=1.1 922	$\phi \pi^0, \phi \rightarrow K^+ K^-$	$(6.6 \pm 0.5) \times 10^{-4}$	—
$2\pi^0$	$(8.25 \pm 0.25) \times 10^{-4}$	923	$2K_S^0 \pi^0$	$< 5.9 \times 10^{-4}$	740
$\pi^+ \pi^- \pi^0$	$(1.47 \pm 0.09) \%$	S=3.0 907	$K^+ K^- \pi^+ \pi^-$	$(2.42 \pm 0.12) \times 10^{-3}$	677
$\rho^+ \pi^-$	$(1.00 \pm 0.06) \%$	764	$\phi(\pi^+ \pi^-)_{S\text{-wave}}, \phi \rightarrow K^+ K^-$	$(2.50 \pm 0.34) \times 10^{-4}$	614
$\rho^0 \pi^0$	$(3.82 \pm 0.29) \times 10^{-3}$	764	$(\phi \rho^0)_{S\text{-wave}}, \phi \rightarrow K^+ K^-$	$(9.3 \pm 1.2) \times 10^{-4}$	250
$\rho^- \pi^+$	$(5.09 \pm 0.34) \times 10^{-3}$	764	$(\phi \rho^0)_{D\text{-wave}}, \phi \rightarrow K^+ K^-$	$(8.2 \pm 2.3) \times 10^{-5}$	—
$\rho(1450)^+ \pi^-, \rho(1450)^+ \rightarrow \pi^+ \pi^0$	$(1.6 \pm 2.0) \times 10^{-5}$	—	$(K^{*0} \bar{K}^{*0})_{S\text{-wave}}, K^{*0} \rightarrow K^\pm \pi^\mp$	$(1.48 \pm 0.30) \times 10^{-4}$	—
$\rho(1450)^0 \pi^0, \rho(1450)^0 \rightarrow \pi^+ \pi^-$	$(4.4 \pm 1.9) \times 10^{-5}$	—			

## Meson Summary Table

$(K^-\pi^+)_{P-wave},$	$(2.6 \pm 0.5) \times 10^{-4}$	—	$\pi^0 e^\pm \mu^\mp$	$LF$	$[hh] < 8.6$	$\times 10^{-5}$	CL=90%	924
$(K^+\pi^-)_{S-wave},$			$\eta e^\pm \mu^\mp$	$LF$	$[hh] < 1.0$	$\times 10^{-4}$	CL=90%	848
$K_1(1270)^+ K^-,$	$(1.8 \pm 0.5) \times 10^{-4}$	—	$\pi^+ \pi^- e^\pm \mu^\mp$	$LF$	$[hh] < 1.5$	$\times 10^{-5}$	CL=90%	911
$K_1(1270)^+ \rightarrow K^{*0} \pi^+$			$\rho^0 e^\pm \mu^\mp$	$LF$	$[hh] < 4.9$	$\times 10^{-5}$	CL=90%	767
$K_1(1270)^+ K^-,$	$(1.14 \pm 0.26) \times 10^{-4}$	—	$\omega e^\pm \mu^\mp$	$LF$	$[hh] < 1.2$	$\times 10^{-4}$	CL=90%	764
$K_1(1270)^+ \rightarrow \rho^0 K^+$			$K^- K^+ e^\pm \mu^\mp$	$LF$	$[hh] < 1.8$	$\times 10^{-4}$	CL=90%	754
$K_1(1270)^- K^+,$	$(2.2 \pm 1.2) \times 10^{-5}$	—	$\phi e^\pm \mu^\mp$	$LF$	$[hh] < 3.4$	$\times 10^{-5}$	CL=90%	648
$K_1(1270)^- \rightarrow \bar{K}^{*0} \pi^-$			$\bar{K}^0 e^\pm \mu^\mp$	$LF$	$[hh] < 1.0$	$\times 10^{-4}$	CL=90%	863
$K_1(1270)^- K^+,$	$(1.45 \pm 0.25) \times 10^{-4}$	—	$K^- \pi^+ e^\pm \mu^\mp$	$LF$	$[hh] < 5.53$	$\times 10^{-4}$	CL=90%	848
$K_1(1270)^- \rightarrow \rho^0 K^-$			$\bar{K}^*(892)^0 e^\pm \mu^\mp$	$LF$	$[hh] < 8.3$	$\times 10^{-5}$	CL=90%	714
$K^*(1410)^+ K^-,$	$(1.02 \pm 0.26) \times 10^{-4}$	—	$2\pi^- 2e^+ + \text{c.c.}$	$L$	$< 1.12$	$\times 10^{-4}$	CL=90%	922
$K^*(1410)^+ \rightarrow K^{*0} \pi^+$			$2\pi^- 2\mu^+ + \text{c.c.}$	$L$	$< 2.9$	$\times 10^{-5}$	CL=90%	894
$K^*(1410)^- K^+,$	$(1.14 \pm 0.25) \times 10^{-4}$	—	$K^- \pi^- 2e^+ + \text{c.c.}$	$L$	$< 2.06$	$\times 10^{-4}$	CL=90%	861
$K^*(1410)^- \rightarrow \bar{K}^{*0} \pi^-$			$K^- \pi^- 2\mu^+ + \text{c.c.}$	$L$	$< 3.9$	$\times 10^{-4}$	CL=90%	829
$2K_S^0 \pi^+ \pi^-$	$(1.24 \pm 0.24) \times 10^{-3}$	673	$2K^- 2e^+ + \text{c.c.}$	$L$	$< 1.52$	$\times 10^{-4}$	CL=90%	791
$K_S^0 K^- 2\pi^+ \pi^-$	$< 1.5 \times 10^{-4}$	CL=90% 595	$2K^- 2\mu^+ + \text{c.c.}$	$L$	$< 9.4$	$\times 10^{-5}$	CL=90%	710
$K^+ K^- \pi^+ \pi^- \pi^0$	$(3.1 \pm 2.0) \times 10^{-3}$	600	$\pi^- \pi^- e^+ \mu^+ + \text{c.c.}$	$L$	$< 7.9$	$\times 10^{-5}$	CL=90%	911
			$K^- \pi^- e^+ \mu^+ + \text{c.c.}$	$L$	$< 2.18$	$\times 10^{-4}$	CL=90%	848
			$2K^- e^+ \mu^+ + \text{c.c.}$	$L$	$< 5.7$	$\times 10^{-5}$	CL=90%	754
			$p e^-$	$L, B$	$[jij] < 1.0$	$\times 10^{-5}$	CL=90%	696
			$\bar{p} e^+$	$L, B$	$[kkk] < 1.1$	$\times 10^{-5}$	CL=90%	696
Other $K\bar{K}X$ modes. They include all decay modes of the $\phi$ , $\eta$ , and $\omega$ .								
$\phi \eta$	$(1.4 \pm 0.5) \times 10^{-4}$	489						
$\phi \omega$	$< 2.1 \times 10^{-3}$	CL=90% 238						
<b>Radiative modes</b>								
$\rho^0 \gamma$	$< 2.4 \times 10^{-4}$	CL=90% 771						
$\omega \gamma$	$< 2.4 \times 10^{-4}$	CL=90% 768						
$\phi \gamma$	$(2.73 \pm 0.35) \times 10^{-5}$	654						
$\bar{K}^*(892)^0 \gamma$	$(3.31 \pm 0.34) \times 10^{-4}$	719						
<b>Doubly Cabibbo suppressed (DC) modes or <math>\Delta C = 2</math> forbidden via mixing (C2M) modes</b>								
$K^+ \ell^- \bar{\nu}_\ell$ via $\bar{D}^0$	$< 2.2 \times 10^{-5}$	CL=90%	—					
$K^+$ or $K^*(892)^+ e^- \bar{\nu}_e$ via $\bar{D}^0$	$< 6 \times 10^{-5}$	CL=90%	—					
$K^+ \pi^-$ DC	$(1.49 \pm 0.07) \times 10^{-4}$	S=2.9 861	—					
$K^+ \pi^-$ via DCS	$(1.33 \pm 0.09) \times 10^{-4}$	—	—					
$K^+ \pi^-$ via $\bar{D}^0$	$< 1.6 \times 10^{-5}$	CL=95% 861	—					
$K_S^0 \pi^+ \pi^-$ in $D^0 \rightarrow \bar{D}^0$	$< 1.8 \times 10^{-4}$	CL=95%	—					
$K^*(892)^+ \pi^-$ , $K^*(892)^+ \rightarrow K_S^0 \pi^+$	$(1.15 \pm_{-0.34}^{+0.60}) \times 10^{-4}$	711	—					
$K_0^*(1430)^+ \pi^-$ , $K_0^*(1430)^+ \rightarrow K_S^0 \pi^+$	$< 1.4 \times 10^{-5}$	—	—					
$K_2^*(1430)^+ \pi^-$ , $K_2^*(1430)^+ \rightarrow K_S^0 \pi^+$	$< 3.4 \times 10^{-5}$	—	—					
$K^+ \pi^- \pi^0$ DC	$(3.13 \pm 0.23) \times 10^{-4}$	844	—					
$K^+ \pi^- \pi^0$ via $\bar{D}^0$	$(7.5 \pm 0.6) \times 10^{-4}$	—	—					
$K^+ \pi^+ 2\pi^-$ DC	$(2.62 \pm 0.11) \times 10^{-4}$	813	—					
$K^+ \pi^+ 2\pi^-$ via $\bar{D}^0$	$< 4 \times 10^{-4}$	CL=90% 812	—					
$\mu^-$ anything via $\bar{D}^0$	$< 4 \times 10^{-4}$	CL=90%	—					
<b><math>\Delta C = 1</math> weak neutral current (C1) modes, Lepton Family number (LF) violating modes, Lepton (L) or Baryon (B) number violating modes</b>								
$\gamma \gamma$	$< 2.2 \times 10^{-6}$	CL=90% 932						
$e^+ e^-$	$< 7.9 \times 10^{-8}$	CL=90% 932						
$\mu^+ \mu^-$	$< 6.2 \times 10^{-9}$	CL=90% 926						
$\pi^0 e^+ e^-$	$< 4.5 \times 10^{-5}$	CL=90% 928						
$\pi^0 \mu^+ \mu^-$	$< 1.8 \times 10^{-4}$	CL=90% 915						
$\eta e^+ e^-$	$< 1.1 \times 10^{-4}$	CL=90% 852						
$\eta \mu^+ \mu^-$	$< 5.3 \times 10^{-4}$	CL=90% 838						
$\pi^+ \pi^- e^+ e^-$	$< 3.73 \times 10^{-4}$	CL=90% 922						
$\rho^0 e^+ e^-$	$< 1.0 \times 10^{-4}$	CL=90% 771						
$\pi^+ \pi^- \mu^+ \mu^-$	$< 5.5 \times 10^{-7}$	CL=90% 894						
$\rho^0 \mu^+ \mu^-$	$< 2.2 \times 10^{-5}$	CL=90% 754						
$\omega e^+ e^-$	$< 1.8 \times 10^{-4}$	CL=90% 768						
$\omega \mu^+ \mu^-$	$< 8.3 \times 10^{-4}$	CL=90% 751						
$K^- K^+ e^+ e^-$	$< 3.15 \times 10^{-4}$	CL=90% 791						
$\phi e^+ e^-$	$< 5.2 \times 10^{-5}$	CL=90% 654						
$K^- K^+ \mu^+ \mu^-$	$< 3.3 \times 10^{-5}$	CL=90% 710						
$\phi \mu^+ \mu^-$	$< 3.1 \times 10^{-5}$	CL=90% 631						
$\bar{K}^0 e^+ e^-$	[zz] $< 1.1 \times 10^{-4}$	CL=90% 866						
$\bar{K}^0 \mu^+ \mu^-$	[zz] $< 2.6 \times 10^{-4}$	CL=90% 852						
$K^- \pi^+ e^+ e^-$	$< 3.85 \times 10^{-4}$	CL=90% 861						
$\bar{K}^*(892)^0 e^+ e^-$	[zz] $< 4.7 \times 10^{-5}$	CL=90% 719						
$K^- \pi^+ \mu^+ \mu^-$	$< 3.59 \times 10^{-4}$	CL=90% 829						
$\bar{K}^*(892)^0 \mu^+ \mu^-$	[zz] $< 2.4 \times 10^{-5}$	CL=90% 700						
$\pi^+ \pi^- \pi^0 \mu^+ \mu^-$	$< 8.1 \times 10^{-4}$	CL=90% 863						
$\mu^\pm e^\mp$	LF [hh] $< 2.6 \times 10^{-7}$	CL=90% 929						

$D^*(2007)^0$

$I(J^P) = \frac{1}{2}(1^-)$   
 $I, J, P$  need confirmation.

Mass  $m = 2006.85 \pm 0.05$  MeV (S = 1.1)  
 $m_{D^{*0}} - m_{D^0} = 142.016 \pm 0.030$  MeV (S = 1.5)  
Full width  $\Gamma < 2.1$  MeV, CL = 90%

$\bar{D}^*(2007)^0$  modes are charge conjugates of modes below.

$D^*(2007)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$D^0 \pi^0$	(64.7 $\pm$ 0.9) %	43
$D^0 \gamma$	(35.3 $\pm$ 0.9) %	137

$D^*(2010)^\pm$

$I(J^P) = \frac{1}{2}(1^-)$   
 $I, J, P$  need confirmation.

Mass  $m = 2010.26 \pm 0.05$  MeV  
 $m_{D^{*}(2010)^+} - m_{D^+} = 140.68 \pm 0.08$  MeV  
 $m_{D^{*}(2010)^+} - m_{D^0} = 145.4257 \pm 0.0017$  MeV  
Full width  $\Gamma = 83.4 \pm 1.8$  keV

$D^*(2010)^-$  modes are charge conjugates of the modes below.

$D^*(2010)^\pm$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$D^0 \pi^+$	(67.7 $\pm$ 0.5) %	39
$D^+ \pi^0$	(30.7 $\pm$ 0.5) %	38
$D^+ \gamma$	(1.6 $\pm$ 0.4) %	136

$D_0^*(2400)^0$

$I(J^P) = \frac{1}{2}(0^+)$

Mass  $m = 2318 \pm 29$  MeV (S = 1.7)  
Full width  $\Gamma = 267 \pm 40$  MeV

$D_0^*(2400)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$D^+ \pi^-$	seen	385

$D_1(2420)^0$

$I(J^P) = \frac{1}{2}(1^+)$   
 $I$  needs confirmation.

Mass  $m = 2420.8 \pm 0.5$  MeV (S = 1.3)  
 $m_{D_1^0} - m_{D^{*+}} = 410.6 \pm 0.5$  (S = 1.3)  
Full width  $\Gamma = 31.7 \pm 2.5$  MeV (S = 3.5)

 **$D^*(2007)^0$** 

$$I(J^P) = \frac{1}{2}(1^-)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 2006.85 \pm 0.05 \text{ MeV} \quad (S = 1.1)$$

$$m_{D^{*0}} - m_{D^0} = 142.016 \pm 0.030 \text{ MeV} \quad (S = 1.5)$$

$$\text{Full width } \Gamma < 2.1 \text{ MeV, CL} = 90\%$$

$\bar{D}^*(2007)^0$  modes are charge conjugates of modes below.

$D^*(2007)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D^0 \pi^0$	(64.7 $\pm$ 0.9) %	43
$D^0 \gamma$	(35.3 $\pm$ 0.9) %	137

 **$D^*(2010)^\pm$** 

$$I(J^P) = \frac{1}{2}(1^-)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 2010.26 \pm 0.05 \text{ MeV}$$

$$m_{D^*(2010)^+} - m_{D^+} = 140.68 \pm 0.08 \text{ MeV}$$

$$m_{D^*(2010)^+} - m_{D^0} = 145.4257 \pm 0.0017 \text{ MeV}$$

$$\text{Full width } \Gamma = 83.4 \pm 1.8 \text{ keV}$$

$D^*(2010)^\mp$  modes are charge conjugates of the modes below.

$D^*(2010)^\pm$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D^0 \pi^+$	(67.7 $\pm$ 0.5) %	39
$D^+ \pi^0$	(30.7 $\pm$ 0.5) %	38
$D^+ \gamma$	(1.6 $\pm$ 0.4) %	136

 **$D_0^*(2400)^0$** 

$$I(J^P) = \frac{1}{2}(0^+)$$

$$\text{Mass } m = 2318 \pm 29 \text{ MeV} \quad (S = 1.7)$$

$$\text{Full width } \Gamma = 267 \pm 40 \text{ MeV}$$

$D_0^*(2400)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D^+ \pi^-$	seen	385

 **$D_1(2420)^0$** 

$$I(J^P) = \frac{1}{2}(1^+)$$

$I$  needs confirmation.

$$\text{Mass } m = 2420.8 \pm 0.5 \text{ MeV} \quad (S = 1.3)$$

$$m_{D_1^0} - m_{D^{*+}} = 410.6 \pm 0.5 \text{ MeV} \quad (S = 1.3)$$

$$\text{Full width } \Gamma = 31.7 \pm 2.5 \text{ MeV} \quad (S = 3.5)$$

# Meson Summary Table

$\overline{D}_1(2420)^0$  modes are charge conjugates of modes below.

$D_1(2420)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D^{*0}(2010)^+ \pi^-$	seen	353
$D^0 \pi^+ \pi^-$	seen	425
$D^+ \pi^-$	not seen	472
$D^{*0} \pi^+ \pi^-$	not seen	279

$D_2^{*}(2460)^0$

$I(J^P) = \frac{1}{2}(2^+)$

$J^P = 2^+$  assignment strongly favored.

Mass  $m = 2460.57 \pm 0.15$  MeV ( $S = 1.1$ )  
 $m_{D_2^{*0}} - m_{D^{*+}} = 590.98 \pm 0.18$  MeV ( $S = 1.1$ )  
 $m_{D_2^{*0}} - m_{D^{*++}} = 450.31 \pm 0.16$  MeV ( $S = 1.1$ )  
Full width  $\Gamma = 47.7 \pm 1.3$  MeV ( $S = 2.0$ )

$\overline{D}_2^{*}(2460)^0$  modes are charge conjugates of modes below.

$D_2^{*}(2460)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D^+ \pi^-$	seen	505
$D^{*0}(2010)^+ \pi^-$	seen	389
$D^0 \pi^+ \pi^-$	not seen	462
$D^{*0} \pi^+ \pi^-$	not seen	324

$D_2^{*}(2460)^{\pm}$

$I(J^P) = \frac{1}{2}(2^+)$

$J^P = 2^+$  assignment strongly favored.

Mass  $m = 2465.4 \pm 1.3$  MeV ( $S = 3.1$ )  
 $m_{D_2^{*}(2460)^{\pm}} - m_{D_2^{*}(2460)^0} = 2.4 \pm 1.7$  MeV  
Full width  $\Gamma = 46.7 \pm 1.2$  MeV

$D_2^{*}(2460)^-$  modes are charge conjugates of modes below.

$D_2^{*}(2460)^{\pm}$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D^0 \pi^+$	seen	513
$D^{*0} \pi^+$	seen	396
$D^+ \pi^+ \pi^-$	not seen	462
$D^{*+} \pi^+ \pi^-$	not seen	326

## CHARMED, STRANGE MESONS ( $C = S = \pm 1$ )

$D_s^+ = c\overline{s}$ ,  $D_s^- = \overline{c}s$ , similarly for  $D_s^{*}$ 's

$D_s^{\pm}$

$I(J^P) = 0(0^-)$

Mass  $m = 1968.27 \pm 0.10$  MeV  
 $m_{D_s^{\pm}} - m_{D^{\pm}} = 98.69 \pm 0.05$  MeV  
Mean life  $\tau = (500 \pm 7) \times 10^{-15}$  s ( $S = 1.3$ )  
 $c\tau = 149.9$   $\mu$ m

### CP-violating decay-rate asymmetries

$A_{CP}(\mu^{\pm}\nu) = (5 \pm 6)\%$   
 $A_{CP}(K^{\pm}K_S^0) = (0.08 \pm 0.26)\%$   
 $A_{CP}(K^+K^-\pi^{\pm}) = (-0.5 \pm 0.9)\%$   
 $A_{CP}(\phi\pi^{\pm}) = (-0.38 \pm 0.27)\%$   
 $A_{CP}(K^{\pm}K_S^0\pi^0) = (-2 \pm 6)\%$   
 $A_{CP}(2K_S^0\pi^{\pm}) = (3 \pm 5)\%$   
 $A_{CP}(K^+K^-\pi^{\pm}\pi^0) = (0.0 \pm 3.0)\%$   
 $A_{CP}(K^{\pm}K_S^0\pi^+\pi^-) = (-6 \pm 5)\%$   
 $A_{CP}(K_S^0K^{\pm}2\pi^{\pm}) = (4.1 \pm 2.8)\%$   
 $A_{CP}(\pi^+\pi^-\pi^{\pm}) = (-0.7 \pm 3.1)\%$   
 $A_{CP}(\pi^{\pm}\eta) = (1.1 \pm 3.1)\%$   
 $A_{CP}(\pi^{\pm}\eta') = (-2.2 \pm 2.3)\%$   
 $A_{CP}(\eta\pi^{\pm}\pi^0) = (-1 \pm 4)\%$   
 $A_{CP}(\eta'\pi^{\pm}\pi^0) = (0 \pm 8)\%$   
 $A_{CP}(K^{\pm}\pi^0) = (-27 \pm 24)\%$   
 $A_{CP}(\overline{K}^0/K^0\pi^{\pm}) = (0.4 \pm 0.5)\%$   
 $A_{CP}(K_S^0\pi^{\pm}) = (3.1 \pm 2.6)\%$  ( $S = 1.7$ )  
 $A_{CP}(K^{\pm}\pi^+\pi^-) = (4 \pm 5)\%$   
 $A_{CP}(K^{\pm}\eta) = (9 \pm 15)\%$   
 $A_{CP}(K^{\pm}\eta'(958)) = (6 \pm 19)\%$

### CP violating asymmetries of P-odd (T-odd) moments

$$A_T(K_S^0K^{\pm}\pi^+\pi^-) = (-14 \pm 8) \times 10^{-3} \text{ [ss]}$$

### $D_s^+ \rightarrow \phi\ell^+\nu_{\ell}$ form factors

$$r_2 = 0.84 \pm 0.11 \quad (S = 2.4)$$
$$r_V = 1.80 \pm 0.08$$
$$\Gamma_L/\Gamma_T = 0.72 \pm 0.18$$

Unless otherwise noted, the branching fractions for modes with a resonance in the final state include all the decay modes of the resonance.  $D_s^-$  modes are charge conjugates of the modes below.

$D_s^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor / Confidence level	$p$ (MeV/c)
Inclusive modes			
$e^+$ semileptonic	[III] ( 6.5 $\pm$ 0.4 ) %		—
$\pi^+$ anything	(119.3 $\pm$ 1.4 ) %		—
$\pi^-$ anything	( 43.2 $\pm$ 0.9 ) %		—
$\pi^0$ anything	(123 $\pm$ 7 ) %		—
$K^-$ anything	( 18.7 $\pm$ 0.5 ) %		—
$K^+$ anything	( 28.9 $\pm$ 0.7 ) %		—
$K_S^0$ anything	( 19.0 $\pm$ 1.1 ) %		—
$\eta$ anything	[nnn] ( 29.9 $\pm$ 2.8 ) %		—
$\omega$ anything	( 6.1 $\pm$ 1.4 ) %		—
$\eta'$ anything	[ooo] ( 10.3 $\pm$ 1.4 ) %	S=1.1	—
$f_0(980)$ anything, $f_0 \rightarrow \pi^+\pi^-$	< 1.3 %	CL=90%	—
$\phi$ anything	( 15.7 $\pm$ 1.0 ) %		—
$K^+K^-$ anything	( 15.8 $\pm$ 0.7 ) %		—
$K_S^0K^+$ anything	( 5.8 $\pm$ 0.5 ) %		—
$K_S^0K^-$ anything	( 1.9 $\pm$ 0.4 ) %		—
$2K_S^0$ anything	( 1.70 $\pm$ 0.32) %		—
$2K^+$ anything	< 2.6 $\times 10^{-3}$	CL=90%	—
$2K^-$ anything	< 6 $\times 10^{-4}$	CL=90%	—

### Leptonic and semileptonic modes

$e^+ \nu_e$	< 8.3 $\times 10^{-5}$	CL=90%	984
$\mu^+ \nu_{\mu}$	( 5.56 $\pm$ 0.25) $\times 10^{-3}$		981
$\tau^+ \nu_{\tau}$	( 5.55 $\pm$ 0.24) %		182
$K^+K^-e^+\nu_e$	—		851
$\phi e^+ \nu_e$	[ppp] ( 2.39 $\pm$ 0.23) %	S=1.8	720
$\eta e^+ \nu_e + \eta'(958) e^+ \nu_e$	[ppp] ( 2.96 $\pm$ 0.29) %		—
$\eta e^+ \nu_e$	[ppp] ( 2.28 $\pm$ 0.24) %		908
$\eta'(958) e^+ \nu_e$	[ppp] ( 6.8 $\pm$ 1.6 ) $\times 10^{-3}$		751
$\omega e^+ \nu_e$	[qqq] < 2.0 $\times 10^{-3}$	CL=90%	829
$K^0 e^+ \nu_e$	( 3.9 $\pm$ 0.9 ) $\times 10^{-3}$		921
$K^*(892)^0 e^+ \nu_e$	[ppp] ( 1.8 $\pm$ 0.4 ) $\times 10^{-3}$		782

### Hadronic modes with a $K\overline{K}$ pair

$K^+K_S^0$	( 1.50 $\pm$ 0.05) %		850
$K^+\overline{K}^0$	( 2.95 $\pm$ 0.14) %		850
$K^+K^-\pi^+$	[tt] ( 5.45 $\pm$ 0.17) %	S=1.2	805
$\phi\pi^+$	[ppp,rrr] ( 4.5 $\pm$ 0.4 ) %		712
$\phi\pi^+, \phi \rightarrow K^+K^-$	( 2.27 $\pm$ 0.08) %		712
$K^+\overline{K}^*(892)^0, \overline{K}^{*0} \rightarrow$	( 2.61 $\pm$ 0.09) %		416
$K^-\pi^+$			
$f_0(980)\pi^+, f_0 \rightarrow K^+K^-$	( 1.15 $\pm$ 0.32) %		732
$f_0(1370)\pi^+, f_0 \rightarrow K^+K^-$	( 7 $\pm$ 5 ) $\times 10^{-4}$		—
$f_0(1710)\pi^+, f_0 \rightarrow K^+K^-$	( 6.7 $\pm$ 2.9 ) $\times 10^{-4}$		198
$K^+\overline{K}_0^*(1430)^0, \overline{K}_0^* \rightarrow$	( 1.9 $\pm$ 0.4 ) $\times 10^{-3}$		218
$K^+K_S^0\pi^0$	( 1.52 $\pm$ 0.22) %		805
$2K_S^0\pi^+$	( 7.7 $\pm$ 0.6 ) $\times 10^{-3}$		802
$K^0\overline{K}^0\pi^+$	—		802
$K^*(892)^+\overline{K}^0$	[ppp] ( 5.4 $\pm$ 1.2 ) %		683
$K^+K^-\pi^+\pi^0$	( 6.3 $\pm$ 0.6 ) %		748
$\phi\rho^+$	[ppp] ( 8.4 $\pm$ 1.9 ) %		401
$K_S^0K^-\pi^+$	( 1.67 $\pm$ 0.10) %		744
$K^*(892)^+\overline{K}^*(892)^0$	[ppp] ( 7.2 $\pm$ 2.6 ) %		416
$K^+K_S^0\pi^+\pi^-$	( 1.03 $\pm$ 0.10) %		744
$K^+K^-\pi^+\pi^-$	( 8.7 $\pm$ 1.5 ) $\times 10^{-3}$		673
$\phi 2\pi^+\pi^-$	[ppp] ( 1.21 $\pm$ 0.16) %		640
$K^+K^-\rho^0\pi^+$ non- $\phi$	< 2.6 $\times 10^{-4}$	CL=90%	249
$\phi\rho^0\pi^+, \phi \rightarrow K^+K^-$	( 6.5 $\pm$ 1.3 ) $\times 10^{-3}$		181
$\phi a_1(1260)^+, \phi \rightarrow$	( 7.5 $\pm$ 1.2 ) $\times 10^{-3}$		†
$K^+K^-, a_1^+ \rightarrow \rho^0\pi^+$			
$K^+K^-\pi^+\pi^-$ nonresonant	( 9 $\pm$ 7 ) $\times 10^{-4}$		673
$2K_S^0 2\pi^+\pi^-$	( 9 $\pm$ 4 ) $\times 10^{-4}$		669

## Meson Summary Table

Hadronic modes without $K$ 's				
$\pi^+ \pi^0$	$< 3.5 \times 10^{-4}$	CL=90%	975	
$2\pi^+ \pi^-$	$(1.09 \pm 0.05) \%$	S=1.1	959	
$\rho^0 \pi^+$	$(2.0 \pm 1.2) \times 10^{-4}$		825	
$\pi^+ (\pi^+ \pi^-)_{S\text{-wave}}$	$[sss] (9.1 \pm 0.4) \times 10^{-3}$		959	
$f_2(1270) \pi^+, f_2 \rightarrow \pi^+ \pi^-$	$(1.10 \pm 0.20) \times 10^{-3}$		559	
$\rho(1450)^0 \pi^+, \rho^0 \rightarrow \pi^+ \pi^-$	$(3.0 \pm 2.0) \times 10^{-4}$		421	
$\pi^+ 2\pi^0$	$(6.5 \pm 1.3) \times 10^{-3}$		960	
$2\pi^+ \pi^- \pi^0$	—		935	
$\eta \pi^+$	$[ppp] (1.70 \pm 0.09) \%$	S=1.1	902	
$\omega \pi^+$	$[ppp] (2.4 \pm 0.6) \times 10^{-3}$		822	
$3\pi^+ 2\pi^-$	$(8.0 \pm 0.8) \times 10^{-3}$		899	
$2\pi^+ \pi^- 2\pi^0$	—		902	
$\eta \rho^+$	$[ppp] (8.9 \pm 0.8) \%$		724	
$\eta \pi^+ \pi^0$	$(9.2 \pm 1.2) \%$		885	
$\omega \pi^+ \pi^0$	$[ppp] (2.8 \pm 0.7) \%$		802	
$3\pi^+ 2\pi^- \pi^0$	$(4.9 \pm 3.2) \%$		856	
$\omega 2\pi^+ \pi^-$	$[ppp] (1.6 \pm 0.5) \%$		766	
$\eta'(958) \pi^+$	$[ooo, ppp] (3.94 \pm 0.25) \%$		743	
$3\pi^+ 2\pi^- 2\pi^0$	—		803	
$\omega \eta \pi^+$	$[ppp] < 2.13 \%$	CL=90%	654	
$\eta'(958) \rho^+$	$[ooo, ppp] (5.8 \pm 1.5) \%$		465	
$\eta'(958) \pi^+ \pi^0$	$(5.6 \pm 0.8) \%$		720	
$\eta'(958) \pi^+ \pi^0$ nonresonant	$< 5.1 \%$	CL=90%	720	
Modes with one or three $K$ 's				
$K^+ \pi^0$	$(6.3 \pm 2.1) \times 10^{-4}$		917	
$K_S^0 \pi^+$	$(1.22 \pm 0.06) \times 10^{-3}$		916	
$K^+ \eta$	$[ppp] (1.77 \pm 0.35) \times 10^{-3}$		835	
$K^+ \omega$	$[ppp] < 2.4 \times 10^{-3}$	CL=90%	741	
$K^+ \eta'(958)$	$[ppp] (1.8 \pm 0.6) \times 10^{-3}$		646	
$K^+ \pi^+ \pi^-$	$(6.6 \pm 0.4) \times 10^{-3}$		900	
$K^+ \rho^0$	$(2.5 \pm 0.4) \times 10^{-3}$		745	
$K^+ \rho(1450)^0, \rho^0 \rightarrow \pi^+ \pi^-$	$(7.0 \pm 2.4) \times 10^{-4}$		—	
$K^*(892)^0 \pi^+, K^{*0} \rightarrow$	$(1.42 \pm 0.24) \times 10^{-3}$		775	
$K^+ \pi^-$	$(1.24 \pm 0.29) \times 10^{-3}$		—	
$K^*(1410)^0 \pi^+, K^{*0} \rightarrow$	$(5.0 \pm 3.5) \times 10^{-4}$		—	
$K^+ \pi^-$	$(1.04 \pm 0.34) \times 10^{-3}$		900	
$K^+ \pi^+ \pi^-$ nonresonant	$(1.00 \pm 0.18) \%$		899	
$K_S^0 2\pi^+ \pi^-$	$(3.0 \pm 1.1) \times 10^{-3}$		870	
$K^+ \omega \pi^0$	$[ppp] < 8.2 \times 10^{-3}$	CL=90%	684	
$K^+ \omega \pi^+ \pi^-$	$[ppp] < 5.4 \times 10^{-3}$	CL=90%	603	
$K^+ \omega \eta$	$[ppp] < 7.9 \times 10^{-3}$	CL=90%	366	
$2K^+ K^-$	$(2.18 \pm 0.21) \times 10^{-4}$		627	
$\phi K^+, \phi \rightarrow K^+ K^-$	$(8.9 \pm 2.0) \times 10^{-5}$		—	
Doubly Cabibbo-suppressed modes				
$2K^+ \pi^-$	$(1.27 \pm 0.13) \times 10^{-4}$		805	
$K^+ K^*(892)^0, K^{*0} \rightarrow$	$(6.0 \pm 3.4) \times 10^{-5}$		—	
$K^+ \pi^-$				
Baryon-antibaryon mode				
$\rho \bar{n}$	$(1.3 \pm 0.4) \times 10^{-3}$		295	
$\Delta C = 1$ weak neutral current (CI) modes, Lepton family number (LF), or Lepton number (L) violating modes				
$\pi^+ e^+ e^-$	$[zz] < 1.3 \times 10^{-5}$	CL=90%	979	
$\pi^+ \phi, \phi \rightarrow e^+ e^-$	$[yy] (6 \pm 4) \times 10^{-6}$		—	
$\pi^+ \mu^+ \mu^-$	$[zz] < 4.1 \times 10^{-7}$	CL=90%	968	
$K^+ e^+ e^-$	$CI < 3.7 \times 10^{-6}$	CL=90%	922	
$K^+ \mu^+ \mu^-$	$CI < 2.1 \times 10^{-5}$	CL=90%	909	
$K^*(892)^+ \mu^+ \mu^-$	$CI < 1.4 \times 10^{-3}$	CL=90%	765	
$\pi^+ e^+ \mu^-$	$LF < 1.2 \times 10^{-5}$	CL=90%	976	
$\pi^+ e^- \mu^+$	$LF < 2.0 \times 10^{-5}$	CL=90%	976	
$K^+ e^+ \mu^-$	$LF < 1.4 \times 10^{-5}$	CL=90%	919	
$K^+ e^- \mu^+$	$LF < 9.7 \times 10^{-6}$	CL=90%	919	
$\pi^- 2e^+$	$L < 4.1 \times 10^{-6}$	CL=90%	979	
$\pi^- 2\mu^+$	$L < 1.2 \times 10^{-7}$	CL=90%	968	
$\pi^- e^+ \mu^+$	$L < 8.4 \times 10^{-6}$	CL=90%	976	
$K^- 2e^+$	$L < 5.2 \times 10^{-6}$	CL=90%	922	
$K^- 2\mu^+$	$L < 1.3 \times 10^{-5}$	CL=90%	909	
$K^- e^+ \mu^+$	$L < 6.1 \times 10^{-6}$	CL=90%	919	
$K^*(892)^- 2\mu^+$	$L < 1.4 \times 10^{-3}$	CL=90%	765	

 **$D_s^{*\pm}$** 

$$I(J^P) = 0(?)^?$$

 $J^P$  is natural, width and decay modes consistent with  $1^-$ .Mass  $m = 2112.1 \pm 0.4$  MeV $m_{D_s^{*+}} - m_{D_s^+} = 143.8 \pm 0.4$  MeVFull width  $\Gamma < 1.9$  MeV, CL = 90% $D_s^{*-}$  modes are charge conjugates of the modes below. **$D_s^{*+}$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D_s^{*+} \gamma$	$(93.5 \pm 0.7) \%$	139
$D_s^{*+} \pi^0$	$(5.8 \pm 0.7) \%$	48
$D_s^{*+} e^+ e^-$	$(6.7 \pm 1.6) \times 10^{-3}$	139

 **$D_{s0}^*(2317)^\pm$** 

$$I(J^P) = 0(0^+)$$

 $J, P$  need confirmation. $J^P$  is natural, low mass consistent with  $0^+$ .Mass  $m = 2317.7 \pm 0.6$  MeV (S = 1.1) $m_{D_{s0}^*(2317)^\pm} - m_{D_s^\pm} = 349.4 \pm 0.6$  MeV (S = 1.1)Full width  $\Gamma < 3.8$  MeV, CL = 95% $D_{s0}^*(2317)^-$  modes are charge conjugates of modes below. **$D_{s0}^*(2317)^\pm$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$D_{s0}^{*+} \pi^0$	seen	298
$D_{s0}^{*+} \pi^0 \pi^0$	not seen	205

 **$D_{s1}(2460)^\pm$** 

$$I(J^P) = 0(1^+)$$

Mass  $m = 2459.5 \pm 0.6$  MeV (S = 1.1) $m_{D_{s1}(2460)^\pm} - m_{D_s^\pm} = 347.3 \pm 0.7$  MeV (S = 1.2) $m_{D_{s1}(2460)^\pm} - m_{D_s^\pm} = 491.2 \pm 0.6$  MeV (S = 1.1)Full width  $\Gamma < 3.5$  MeV, CL = 95% $D_{s1}(2460)^-$  modes are charge conjugates of the modes below. **$D_{s1}(2460)^+$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$\rho$ (MeV/c)
$D_{s1}^{*+} \pi^0$	$(48 \pm 11) \%$		297
$D_{s1}^{*+} \gamma$	$(18 \pm 4) \%$		442
$D_{s1}^{*+} \pi^+ \pi^-$	$(4.3 \pm 1.3) \%$	S=1.1	363
$D_{s1}^{*+} \gamma$	$< 8 \%$	CL=90%	323
$D_{s0}^*(2317)^+ \gamma$	$(3.7 \pm 5.0_{2.4}) \%$		138

 **$D_{s1}(2536)^\pm$** 

$$I(J^P) = 0(1^+)$$

 $J, P$  need confirmation.Mass  $m = 2535.10 \pm 0.06$  MeVFull width  $\Gamma = 0.92 \pm 0.05$  MeV $D_{s1}(2536)^-$  modes are charge conjugates of the modes below. **$D_{s1}(2536)^+$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$D^*(2010)^+ K^0$	$0.85 \pm 0.12$		149
$(D^*(2010)^+ K^0)_{S\text{-wave}}$	$0.61 \pm 0.09$		149
$D^+ \pi^- K^+$	$0.028 \pm 0.005$		176
$D^*(2007)^0 K^+$	<b>DEFINED AS 1</b>		167
$D^+ K^0$	$< 0.34$	90%	381
$D^0 K^+$	$< 0.12$	90%	391
$D_s^{*+} \gamma$	possibly seen		388
$D_s^{*+} \pi^+ \pi^-$	seen		437

 **$D_{s2}^*(2573)$** 

$$I(J^P) = 0(2^+)$$

 $J^P$  is natural, width and decay modes consistent with  $2^+$ .Mass  $m = 2569.1 \pm 0.8$  MeV (S = 2.4)Full width  $\Gamma = 16.9 \pm 0.8$  MeV

# Meson Summary Table

$D_{s2}^*(2573)^-$  modes are charge conjugates of the modes below.

$D_{s2}^*(2573)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$D^0 K^+$	seen	431
$D^*(2007)^0 K^+$	not seen	238

**$D_{s1}^*(2700)^\pm$**

$$I(J^P) = 0(1^-)$$

Mass  $m = 2708.3^{+4.0}_{-3.4}$  MeV

Full width  $\Gamma = 120 \pm 11$  MeV

## BOTTOM MESONS ( $B = \pm 1$ )

$$B^+ = u\bar{b}, B^0 = d\bar{b}, \bar{B}^0 = \bar{d}b, B^- = \bar{u}b, \text{ similarly for } B^{*+}\text{'s}$$

### B-particle organization

Many measurements of  $B$  decays involve admixtures of  $B$  hadrons. Previously we arbitrarily included such admixtures in the  $B^\pm$  section, but because of their importance we have created two new sections: “ $B^\pm/B^0$  Admixture” for  $\Upsilon(4S)$  results and “ $B^\pm/B^0/B_s^0/b$ -baryon Admixture” for results at higher energies. Most inclusive decay branching fractions and  $\chi_b$  at high energy are found in the Admixture sections.  $B^0$ - $\bar{B}^0$  mixing data are found in the  $B^0$  section, while  $B_s^0$ - $\bar{B}_s^0$  mixing data and  $B^0$ - $\bar{B}$  mixing data for a  $B^0/B_s^0$  admixture are found in the  $B_s^0$  section.  $CP$ -violation data are found in the  $B^\pm$ ,  $B^0$ , and  $B^\pm B^0$  Admixture sections.  $b$ -baryons are found near the end of the Baryon section.

The organization of the  $B$  sections is now as follows, where bullets indicate particle sections and brackets indicate reviews.

- $B^\pm$   
mass, mean life,  $CP$  violation, branching fractions
- $B^0$   
mass, mean life,  $B^0$ - $\bar{B}^0$  mixing,  $CP$  violation, branching fractions
- $B^\pm/B^0$  Admixtures  
 $CP$  violation, branching fractions
- $B^\pm/B^0/B_s^0/b$ -baryon Admixtures  
mean life, production fractions, branching fractions
- $B^*$   
mass
- $B_1(5721)^+$   
mass
- $B_1(5721)^0$   
mass
- $B_2^*(5747)^+$   
mass
- $B_2^*(5747)^0$   
mass
- $B_J^*(5970)^+$   
mass
- $B_J^*(5970)^0$   
mass
- $B_s^0$   
mass, mean life,  $B_s^0$ - $\bar{B}_s^0$  mixing,  $CP$  violation, branching fractions
- $B_s^*$   
mass
- $B_{s1}(5830)^0$

mass

•  $B_{s2}^*(5840)^0$

mass

•  $B_c^\pm$

mass, mean life, branching fractions

At the end of Baryon Listings:

•  $\Lambda_b$

mass, mean life, branching fractions

•  $\Lambda_b(5912)^0$

mass, mean life

•  $\Lambda_b(5920)^0$

mass, mean life

•  $\Sigma_b$

mass

•  $\Sigma_b^*$

mass

•  $\Xi_b^0, \Xi_b^-$

mass, mean life, branching fractions

•  $\Xi_b'(5935)^-$

mass

•  $\Xi_b(5945)^0$

mass

•  $\Xi_b^*(5955)^-$

mass

•  $\Omega_b^-$

mass, branching fractions

•  $b$ -baryon Admixture

mean life, branching fractions

**$B^\pm$**

$$I(J^P) = \frac{1}{2}(0^-)$$

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

Mass  $m_{B^\pm} = 5279.31 \pm 0.15$  MeV ( $S = 1.1$ )

Mean life  $\tau_{B^\pm} = (1.638 \pm 0.004) \times 10^{-12}$  s

$c\tau = 491.1 \mu\text{m}$

### CP violation

$$\begin{aligned} A_{CP}(B^+ \rightarrow J/\psi(1S)K^+) &= 0.003 \pm 0.006 \quad (S = 1.8) \\ A_{CP}(B^+ \rightarrow J/\psi(1S)\pi^+) &= (0.1 \pm 2.8) \times 10^{-2} \quad (S = 1.2) \\ A_{CP}(B^+ \rightarrow J/\psi\rho^+) &= -0.11 \pm 0.14 \\ A_{CP}(B^+ \rightarrow J/\psi K^*(892)^+) &= -0.048 \pm 0.033 \\ A_{CP}(B^+ \rightarrow \eta_c K^+) &= 0.01 \pm 0.07 \quad (S = 2.2) \\ A_{CP}(B^+ \rightarrow \psi(2S)\pi^+) &= 0.03 \pm 0.06 \\ A_{CP}(B^+ \rightarrow \psi(2S)K^+) &= 0.012 \pm 0.020 \quad (S = 1.5) \\ A_{CP}(B^+ \rightarrow \psi(2S)K^*(892)^+) &= 0.08 \pm 0.21 \\ A_{CP}(B^+ \rightarrow \chi_{c1}(1P)\pi^+) &= 0.07 \pm 0.18 \\ A_{CP}(B^+ \rightarrow \chi_{c0}K^+) &= -0.20 \pm 0.18 \quad (S = 1.5) \\ A_{CP}(B^+ \rightarrow \chi_{c1}K^+) &= -0.009 \pm 0.033 \\ A_{CP}(B^+ \rightarrow \chi_{c1}K^*(892)^+) &= 0.5 \pm 0.5 \\ A_{CP}(B^+ \rightarrow \bar{D}^0\pi^+) &= -0.007 \pm 0.007 \\ A_{CP}(B^+ \rightarrow D_{CP(+1)}\pi^+) &= 0.035 \pm 0.024 \\ A_{CP}(B^+ \rightarrow D_{CP(-1)}\pi^+) &= 0.017 \pm 0.026 \\ A_{CP}([K^\mp\pi^\pm\pi^+\pi^-]_D\pi^+) &= 0.13 \pm 0.10 \\ A_{CP}(B^+ \rightarrow \bar{D}^0K^+) &= 0.007 \pm 0.025 \quad (S = 1.5) \\ A_{CP}([K^\mp\pi^\pm\pi^+\pi^-]_D K^+) &= -0.42 \pm 0.22 \\ r_B(B^+ \rightarrow D^0 K^+) &= 0.095 \pm 0.008 \\ \delta_B(B^+ \rightarrow D^0 K^+) &= (123 \pm 10)^\circ \\ r_B(B^+ \rightarrow \bar{D}^0 K^{*+}) &= 0.17 \pm 0.11 \quad (S = 2.3) \\ \delta_B(B^+ \rightarrow D^0 K^{*+}) &= (155 \pm 70)^\circ \quad (S = 2.0) \\ A_{CP}(B^+ \rightarrow [K^-\pi^+]_D K^+) &= -0.58 \pm 0.21 \\ A_{CP}(B^+ \rightarrow [K^-\pi^+\pi^0]_D K^+) &= 0.07 \pm 0.30 \quad (S = 1.5) \\ A_{CP}(B^+ \rightarrow [K^+K^-\pi^0]_D K^+) &= 0.30 \pm 0.20 \\ A_{CP}(B^+ \rightarrow [\pi^+\pi^-\pi^0]_D K^+) &= 0.05 \pm 0.09 \\ A_{CP}(B^+ \rightarrow [K^-\pi^+]_D K^*(892)^+) &= -0.3 \pm 0.5 \\ A_{CP}(B^+ \rightarrow [K^-\pi^+]_D \pi^+) &= 0.00 \pm 0.09 \end{aligned}$$

## Meson Summary Table

$A_{CP}(B^+ \rightarrow [K^- \pi^+ \pi^0]_D \pi^+) = 0.35 \pm 0.16$	$A_{CP}(B^+ \rightarrow K^{*+} \pi^+ \pi^-) = 0.07 \pm 0.08$
$A_{CP}(B^+ \rightarrow [K^+ K^- \pi^0]_D \pi^+) = -0.03 \pm 0.04$	$A_{CP}(B^+ \rightarrow \rho^0 K^*(892)^+) = 0.31 \pm 0.13$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D \pi^+) = -0.016 \pm 0.020$	$A_{CP}(B^+ \rightarrow K^*(892)^+ f_0(980)) = -0.15 \pm 0.12$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\pi)} \pi^+) = -0.09 \pm 0.27$	$A_{CP}(B^+ \rightarrow a_1^+ K^0) = 0.12 \pm 0.11$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\gamma)} \pi^+) = -0.7 \pm 0.6$	$A_{CP}(B^+ \rightarrow b_1^+ K^0) = -0.03 \pm 0.15$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\pi)} K^+) = 0.8 \pm 0.4$	$A_{CP}(B^+ \rightarrow K^*(892)^0 \rho^+) = -0.01 \pm 0.16$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\gamma)} K^+) = 0.4 \pm 1.0$	$A_{CP}(B^+ \rightarrow b_1^+ K^+) = -0.46 \pm 0.20$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D K^+) = -0.02 \pm 0.15$	$A_{CP}(B^+ \rightarrow K^0 K^+) = 0.04 \pm 0.14$
$A_{CP}(B^+ \rightarrow [K_S^0 K^+ \pi^-]_D K^+) = 0.04 \pm 0.09$	$A_{CP}(B^+ \rightarrow K_S^0 K^+) = -0.21 \pm 0.14$
$A_{CP}(B^+ \rightarrow [K_S^0 K^- \pi^+]_D K^+) = 0.23 \pm 0.13$	$A_{CP}(B^+ \rightarrow K^+ K_S^0 K_S^0) = 0.04^{+0.04}_{-0.05}$
$A_{CP}(B^+ \rightarrow [K_S^0 K^- \pi^+]_D \pi^+) = -0.052 \pm 0.034$	<b><math>A_{CP}(B^+ \rightarrow K^+ K^- \pi^+) = -0.118 \pm 0.022</math></b>
$A_{CP}(B^+ \rightarrow [K_S^0 K^+ \pi^-]_D \pi^+) = -0.025 \pm 0.026$	<b><math>A_{CP}(B^+ \rightarrow K^+ K^- K^+) = -0.033 \pm 0.008</math></b>
$A_{CP}(B^+ \rightarrow [K^*(892)^- K^+]_D K^+) = 0.03 \pm 0.11$	$A_{CP}(B^+ \rightarrow \phi K^+) = 0.024 \pm 0.028 \quad (S = 2.3)$
$A_{CP}(B^+ \rightarrow [K^*(892)^+ K^-]_D K^+) = 0.34 \pm 0.21$	$A_{CP}(B^+ \rightarrow X_0(1550) K^+) = -0.04 \pm 0.07$
$A_{CP}(B^+ \rightarrow [K^*(892)^+ K^-]_D \pi^+) = -0.05 \pm 0.05$	$A_{CP}(B^+ \rightarrow K^{*+} K^+ K^-) = 0.11 \pm 0.09$
$A_{CP}(B^+ \rightarrow [K^*(892)^- K^+]_D \pi^+) = -0.012 \pm 0.030$	$A_{CP}(B^+ \rightarrow \phi K^*(892)^+) = -0.01 \pm 0.08$
<b><math>A_{CP}(B^+ \rightarrow D_{CP(+1)} K^+) = 0.170 \pm 0.033 \quad (S = 1.2)</math></b>	$A_{CP}(B^+ \rightarrow \phi(K\pi)_0^{*+}) = 0.04 \pm 0.16$
$A_{ADS}(B^+ \rightarrow D K^+) = -0.52 \pm 0.15$	$A_{CP}(B^+ \rightarrow \phi K_1(1270)^+) = 0.15 \pm 0.20$
$A_{ADS}(B^+ \rightarrow D \pi^+) = 0.14 \pm 0.06$	$A_{CP}(B^+ \rightarrow \phi K_2^*(1430)^+) = -0.23 \pm 0.20$
$A_{ADS}(B^+ \rightarrow [K^- \pi^+]_D K^+ \pi^- \pi^+) = -0.33 \pm 0.35$	$A_{CP}(B^+ \rightarrow K^+ \phi \phi) = -0.10 \pm 0.08$
$A_{ADS}(B^+ \rightarrow [K^- \pi^+]_D \pi^+ \pi^- \pi^+) = -0.01 \pm 0.09$	$A_{CP}(B^+ \rightarrow K^+[\phi \phi]_{\eta_c}) = 0.09 \pm 0.10$
$A_{CP}(B^+ \rightarrow D_{CP(-1)} K^+) = -0.10 \pm 0.07$	$A_{CP}(B^+ \rightarrow K^*(892)^+ \gamma) = 0.018 \pm 0.029$
$A_{CP}(B^+ \rightarrow [K^+ K^-]_D K^+ \pi^- \pi^+) = -0.04 \pm 0.06$	$A_{CP}(B^+ \rightarrow \eta K^+ \gamma) = -0.12 \pm 0.07$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^-]_D K^+ \pi^- \pi^+) = -0.05 \pm 0.10$	$A_{CP}(B^+ \rightarrow \phi K^+ \gamma) = -0.13 \pm 0.11 \quad (S = 1.1)$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_D K^+ \pi^- \pi^+) = 0.013 \pm 0.023$	$A_{CP}(B^+ \rightarrow \rho^+ \gamma) = -0.11 \pm 0.33$
$A_{CP}(B^+ \rightarrow [K^+ K^-]_D \pi^+ \pi^- \pi^+) = -0.019 \pm 0.015$	$A_{CP}(B^+ \rightarrow \pi^+ \pi^0) = 0.03 \pm 0.04$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^-]_D \pi^+ \pi^- \pi^+) = -0.013 \pm 0.019$	<b><math>A_{CP}(B^+ \rightarrow \pi^+ \pi^- \pi^+) = 0.057 \pm 0.013</math></b>
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_D \pi^+ \pi^- \pi^+) = -0.002 \pm 0.011$	$A_{CP}(B^+ \rightarrow \rho^0 \pi^+) = 0.18^{+0.09}_{-0.17}$
$A_{CP}(B^+ \rightarrow \bar{D}^{*0} \pi^+) = -0.014 \pm 0.015$	$A_{CP}(B^+ \rightarrow f_2(1270) \pi^+) = 0.41 \pm 0.30$
$A_{CP}(B^+ \rightarrow (D_{CP(+1)}^*)^0 \pi^+) = -0.02 \pm 0.05$	$A_{CP}(B^+ \rightarrow \rho^0(1450) \pi^+) = -0.1^{+0.4}_{-0.5}$
$A_{CP}(B^+ \rightarrow (D_{CP(-1)}^*)^0 \pi^+) = -0.09 \pm 0.05$	<b><math>A_{CP}(B^+ \rightarrow f_0(1370) \pi^+) = 0.72 \pm 0.22</math></b>
$A_{CP}(B^+ \rightarrow D^{*0} K^+) = -0.07 \pm 0.04$	$A_{CP}(B^+ \rightarrow \pi^+ \pi^- \pi^+ \text{ nonresonant}) = -0.14^{+0.23}_{-0.16}$
$r_B^*(B^+ \rightarrow D^{*0} K^+) = 0.114^{+0.023}_{-0.040} \quad (S = 1.2)$	$A_{CP}(B^+ \rightarrow \rho^+ \pi^0) = 0.02 \pm 0.11$
$\delta_B^*(B^+ \rightarrow D^{*0} K^+) = (310^{+22}_{-28})^\circ \quad (S = 1.3)$	$A_{CP}(B^+ \rightarrow \rho^+ \rho^0) = -0.05 \pm 0.05$
$A_{CP}(B^+ \rightarrow D_{CP(+1)}^{*0} K^+) = -0.12 \pm 0.08$	$A_{CP}(B^+ \rightarrow \omega \pi^+) = -0.04 \pm 0.06$
$A_{CP}(B^+ \rightarrow D_{CP(-1)}^* K^+) = 0.07 \pm 0.10$	$A_{CP}(B^+ \rightarrow \omega \rho^+) = -0.20 \pm 0.09$
$A_{CP}(B^+ \rightarrow D_{CP(+1)} K^*(892)^+) = 0.09 \pm 0.14$	$A_{CP}(B^+ \rightarrow \eta \pi^+) = -0.14 \pm 0.07 \quad (S = 1.4)$
$A_{CP}(B^+ \rightarrow D_{CP(-1)} K^*(892)^+) = -0.23 \pm 0.22$	$A_{CP}(B^+ \rightarrow \eta \rho^+) = 0.11 \pm 0.11$
$A_{CP}(B^+ \rightarrow D_S^+ \phi) = 0.0 \pm 0.4$	$A_{CP}(B^+ \rightarrow \eta' \pi^+) = 0.06 \pm 0.16$
$A_{CP}(B^+ \rightarrow D^{*+} \bar{D}^{*0}) = -0.15 \pm 0.11$	$A_{CP}(B^+ \rightarrow \eta' \rho^+) = 0.26 \pm 0.17$
$A_{CP}(B^+ \rightarrow D^{*+} \bar{D}^0) = -0.06 \pm 0.13$	$A_{CP}(B^+ \rightarrow b_1^0 \pi^+) = 0.05 \pm 0.16$
$A_{CP}(B^+ \rightarrow D^+ \bar{D}^{*0}) = 0.13 \pm 0.18$	$A_{CP}(B^+ \rightarrow \rho \bar{\rho} \pi^+) = 0.00 \pm 0.04$
$A_{CP}(B^+ \rightarrow D^+ \bar{D}^0) = -0.03 \pm 0.07$	$A_{CP}(B^+ \rightarrow \rho \bar{\rho} K^+) = 0.00 \pm 0.04 \quad (S = 2.2)$
$A_{CP}(B^+ \rightarrow K_S^0 \pi^+) = -0.017 \pm 0.016$	$A_{CP}(B^+ \rightarrow \rho \bar{\rho} K^*(892)^+) = 0.21 \pm 0.16 \quad (S = 1.4)$
$A_{CP}(B^+ \rightarrow K^+ \pi^0) = 0.037 \pm 0.021$	$A_{CP}(B^+ \rightarrow \rho \bar{\rho} \gamma) = 0.17 \pm 0.17$
$A_{CP}(B^+ \rightarrow \eta' K^+) = 0.004 \pm 0.011$	$A_{CP}(B^+ \rightarrow \rho \bar{\rho} \pi^0) = 0.01 \pm 0.17$
$A_{CP}(B^+ \rightarrow \eta' K^*(892)^+) = -0.26 \pm 0.27$	$A_{CP}(B^+ \rightarrow K^+ \ell^+ \ell^-) = -0.02 \pm 0.08$
$A_{CP}(B^+ \rightarrow \eta' K_0^*(1430)^+) = 0.06 \pm 0.20$	$A_{CP}(B^+ \rightarrow K^+ e^+ e^-) = 0.14 \pm 0.14$
$A_{CP}(B^+ \rightarrow \eta' K_2^*(1430)^+) = 0.15 \pm 0.13$	$A_{CP}(B^+ \rightarrow K^+ \mu^+ \mu^-) = 0.011 \pm 0.017$
<b><math>A_{CP}(B^+ \rightarrow \eta K^+) = -0.37 \pm 0.08</math></b>	$A_{CP}(B^+ \rightarrow \pi^+ \mu^+ \mu^-) = -0.11 \pm 0.12$
$A_{CP}(B^+ \rightarrow \eta K^*(892)^+) = 0.02 \pm 0.06$	$A_{CP}(B^+ \rightarrow K^{*+} \ell^+ \ell^-) = -0.09 \pm 0.14$
$A_{CP}(B^+ \rightarrow \eta K_0^*(1430)^+) = 0.05 \pm 0.13$	$A_{CP}(B^+ \rightarrow K^* e^+ e^-) = -0.14 \pm 0.23$
$A_{CP}(B^+ \rightarrow \eta K_2^*(1430)^+) = -0.45 \pm 0.30$	$A_{CP}(B^+ \rightarrow K^* \mu^+ \mu^-) = -0.12 \pm 0.24$
$A_{CP}(B^+ \rightarrow \omega K^+) = -0.02 \pm 0.04$	<b><math>\gamma(B^+ \rightarrow D^{(*)0} K^{(*)+}) = (70 \pm 9)^\circ</math></b>
$A_{CP}(B^+ \rightarrow \omega K^{*+}) = 0.29 \pm 0.35$	$\gamma(B^+ \rightarrow D K^+ \pi^- \pi^+, D \pi^+ \pi^- \pi^+) = (74 \pm 20)^\circ$
$A_{CP}(B^+ \rightarrow \omega(K\pi)_0^{*+}) = -0.10 \pm 0.09$	
$A_{CP}(B^+ \rightarrow \omega K_2^*(1430)^+) = 0.14 \pm 0.15$	
$A_{CP}(B^+ \rightarrow K^{*0} \pi^+) = -0.04 \pm 0.09 \quad (S = 2.1)$	
$A_{CP}(B^+ \rightarrow K^*(892)^+ \pi^0) = -0.06 \pm 0.24$	
<b><math>A_{CP}(B^+ \rightarrow K^+ \pi^- \pi^+) = 0.027 \pm 0.008</math></b>	
$A_{CP}(B^+ \rightarrow K^+ K^- K^+ \text{ nonresonant}) = 0.06 \pm 0.05$	
$A_{CP}(B^+ \rightarrow f(980)^0 K^+) = -0.08 \pm 0.09$	
<b><math>A_{CP}(B^+ \rightarrow f_2(1270) K^+) = -0.68^{+0.19}_{-0.17}</math></b>	
$A_{CP}(B^+ \rightarrow f_0(1500) K^+) = 0.28 \pm 0.30$	
$A_{CP}(B^+ \rightarrow f_2'(1525)^0 K^+) = -0.08^{+0.05}_{-0.04}$	
<b><math>A_{CP}(B^+ \rightarrow \rho^0 K^+) = 0.37 \pm 0.10</math></b>	
$A_{CP}(B^+ \rightarrow K_0^*(1430)^0 \pi^+) = 0.055 \pm 0.033$	
$A_{CP}(B^+ \rightarrow K_2^*(1430)^0 \pi^+) = 0.05^{+0.29}_{-0.24}$	
$A_{CP}(B^+ \rightarrow K^+ \pi^0 \pi^0) = -0.06 \pm 0.07$	
$A_{CP}(B^+ \rightarrow K^0 \rho^+) = -0.12 \pm 0.17$	

# Meson Summary Table

$B^-$  modes are charge conjugates of the modes below. Modes which do not identify the charge state of the  $B$  are listed in the  $B^\pm/B^0$  ADMIXTURE section.

The branching fractions listed below assume 50%  $B^0\bar{B}^0$  and 50%  $B^+B^-$  production at the  $\Upsilon(4S)$ . We have attempted to bring older measurements up to date by rescaling their assumed  $\Upsilon(4S)$  production ratio to 50:50 and their assumed  $D, D_s, D^*$ , and  $\psi$  branching ratios to current values whenever this would affect our averages and best limits significantly.

Indentation is used to indicate a subchannel of a previous reaction. All resonant subchannels have been corrected for resonance branching fractions to the final state so the sum of the subchannel branching fractions can exceed that of the final state.

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm$  anything, the values usually are multiplicities, not branching fractions. They can be greater than one.

B <sup>+</sup> DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level (MeV/c)	$\rho$
<b>Semileptonic and leptonic modes</b>			
$\ell^+ \nu_\ell$ anything	[ttt] ( 10.99 $\pm$ 0.28 ) %	—	—
$e^+ \nu_e X_C$	( 10.8 $\pm$ 0.4 ) %	—	—
$D \ell^+ \nu_\ell$ anything	( 9.8 $\pm$ 0.7 ) %	—	—
$\bar{D}^0 \ell^+ \nu_\ell$	[ttt] ( 2.27 $\pm$ 0.11 ) %	2310	—
$\bar{D}^0 \tau^+ \nu_\tau$	( 7.7 $\pm$ 2.5 ) $\times 10^{-3}$	1911	—
$\bar{D}^*(2007)^0 \ell^+ \nu_\ell$	[ttt] ( 5.69 $\pm$ 0.19 ) %	2258	—
$\bar{D}^*(2007)^0 \tau^+ \nu_\tau$	( 1.88 $\pm$ 0.20 ) %	1839	—
$D^- \pi^+ \ell^+ \nu_\ell$	( 4.2 $\pm$ 0.5 ) $\times 10^{-3}$	2306	—
$\bar{D}_0^*(2420)^0 \ell^+ \nu_\ell, \bar{D}_0^{*0} \rightarrow$	( 2.5 $\pm$ 0.5 ) $\times 10^{-3}$	—	—
$\bar{D}_2^{*0}(2460)^0 \ell^+ \nu_\ell, \bar{D}_2^{*0} \rightarrow$	( 1.53 $\pm$ 0.16 ) $\times 10^{-3}$	2065	—
$D^{(*)-} n \pi^+ \ell^+ \nu_\ell (n \geq 1)$	( 1.87 $\pm$ 0.26 ) %	—	—
$D^{*-} \pi^+ \ell^+ \nu_\ell$	( 6.1 $\pm$ 0.6 ) $\times 10^{-3}$	2254	—
$\bar{D}_1(2420)^0 \ell^+ \nu_\ell, \bar{D}_1^0 \rightarrow$	( 3.03 $\pm$ 0.20 ) $\times 10^{-3}$	2084	—
$\bar{D}_1^{*-} \pi^+ \ell^+ \nu_\ell, \bar{D}_1^{*0} \rightarrow$	( 2.7 $\pm$ 0.6 ) $\times 10^{-3}$	—	—
$\bar{D}_2^{*-} \pi^+ \ell^+ \nu_\ell, \bar{D}_2^{*0}(2460)^0 \ell^+ \nu_\ell,$	( 1.01 $\pm$ 0.24 ) $\times 10^{-3}$	S=2.0 2065	—
$\bar{D}_2^{*0} \rightarrow D^{*-} \pi^+$	—	—	—
$\bar{D}^0 \pi^+ \pi^- \ell^+ \nu_\ell$	( 1.6 $\pm$ 0.4 ) $\times 10^{-3}$	2301	—
$\bar{D}^{*0} \pi^+ \pi^- \ell^+ \nu_\ell$	( 8 $\pm$ 5 ) $\times 10^{-4}$	2248	—
$D_s^{(*)-} K^+ \ell^+ \nu_\ell$	( 6.1 $\pm$ 1.0 ) $\times 10^{-4}$	—	—
$D_s^- K^+ \ell^+ \nu_\ell$	( 3.0 $\pm$ 1.4 ) $\times 10^{-4}$	2242	—
$D_s^{*-} K^+ \ell^+ \nu_\ell$	( 2.9 $\pm$ 1.9 ) $\times 10^{-4}$	2185	—
$\pi^0 \ell^+ \nu_\ell$	( 7.80 $\pm$ 0.27 ) $\times 10^{-5}$	2638	—
$\eta \ell^+ \nu_\ell$	( 3.8 $\pm$ 0.6 ) $\times 10^{-5}$	2611	—
$\eta' \ell^+ \nu_\ell$	( 2.3 $\pm$ 0.8 ) $\times 10^{-5}$	2553	—
$\omega \ell^+ \nu_\ell$	[ttt] ( 1.19 $\pm$ 0.09 ) $\times 10^{-4}$	2582	—
$\rho^0 \ell^+ \nu_\ell$	[ttt] ( 1.58 $\pm$ 0.11 ) $\times 10^{-4}$	2583	—
$\rho \bar{p} \ell^+ \nu_\ell$	( 5.8 $\pm$ 2.6 ) $\times 10^{-6}$	2467	—
$\rho \bar{p} \mu^+ \nu_\mu$	< 8.5 $\times 10^{-6}$	CL=90% 2446	—
$\rho \bar{p} e^+ \nu_e$	( 8.2 $\pm$ 4.0 ) $\times 10^{-6}$	2467	—
$e^+ \nu_e$	< 9.8 $\times 10^{-7}$	CL=90% 2640	—
$\mu^+ \nu_\mu$	< 1.0 $\times 10^{-6}$	CL=90% 2639	—
$\tau^+ \nu_\tau$	( 1.09 $\pm$ 0.24 ) $\times 10^{-4}$	S=1.2 2341	—
$\ell^+ \nu_\ell \gamma$	< 3.5 $\times 10^{-6}$	CL=90% 2640	—
$e^+ \nu_e \gamma$	< 6.1 $\times 10^{-6}$	CL=90% 2640	—
$\mu^+ \nu_\mu \gamma$	< 3.4 $\times 10^{-6}$	CL=90% 2639	—
<b>Inclusive modes</b>			
$D^0 X$	( 8.6 $\pm$ 0.7 ) %	—	—
$\bar{D}^0 X$	( 79 $\pm$ 4 ) %	—	—
$D^+ X$	( 2.5 $\pm$ 0.5 ) %	—	—
$D^- X$	( 9.9 $\pm$ 1.2 ) %	—	—
$D_s^+ X$	( 7.9 $\pm$ 1.4 ) %	—	—
$D_s^- X$	( 1.10 $\pm$ 0.40 ) %	—	—
$\Lambda_c^+ X$	( 2.1 $\pm$ 0.9 ) %	—	—
$\bar{\Lambda}_c^- X$	( 2.8 $\pm$ 1.1 ) %	—	—
$\bar{c} X$	( 97 $\pm$ 4 ) %	—	—
$c X$	( 23.4 $\pm$ 2.2 ) %	—	—
$c/\bar{c} X$	( 120 $\pm$ 6 ) %	—	—

## D, D\*, or D<sub>s</sub> modes

$\bar{D}^0 \pi^+$	( 4.80 $\pm$ 0.15 ) $\times 10^{-3}$	2308	—
$D_{CP(+1)} \pi^+$	[uuu] ( 2.19 $\pm$ 0.24 ) $\times 10^{-3}$	—	—
$D_{CP(-1)} \pi^+$	[uuu] ( 2.1 $\pm$ 0.4 ) $\times 10^{-3}$	—	—
$\bar{D}^0 \rho^+$	( 1.34 $\pm$ 0.18 ) %	2237	—
$\bar{D}^0 K^+$	( 3.69 $\pm$ 0.17 ) $\times 10^{-4}$	2281	—
$D_{CP(+1)} K^+$	[uuu] ( 1.91 $\pm$ 0.14 ) $\times 10^{-4}$	—	—
$D_{CP(-1)} K^+$	[uuu] ( 1.99 $\pm$ 0.19 ) $\times 10^{-4}$	—	—
$[K^- \pi^+]_D K^+$	[vvv] < 2.8 $\times 10^{-7}$	CL=90% —	—
$[K^+ \pi^-]_D K^+$	[vvv] < 1.8 $\times 10^{-5}$	CL=90% —	—
$[K^- \pi^+ \pi^0]_D K^+$	seen	—	—
$[K^+ \pi^- \pi^0]_D K^+$	seen	—	—
$[K^- \pi^+ \pi^+ \pi^-]_D K^+$	seen	—	—
$[K^+ \pi^- \pi^+ \pi^-]_D K^+$	seen	—	—
$[K^- \pi^+]_D \pi^+$	[vvv] ( 6.3 $\pm$ 1.1 ) $\times 10^{-7}$	—	—
$[K^+ \pi^-]_D \pi^+$	( 1.68 $\pm$ 0.31 ) $\times 10^{-4}$	—	—
$[K^- \pi^+ \pi^0]_D \pi^+$	seen	—	—
$[K^+ \pi^- \pi^0]_D \pi^+$	seen	—	—
$[K^- \pi^+ \pi^+ \pi^-]_D \pi^+$	seen	—	—
$[K^+ \pi^- \pi^+ \pi^-]_D \pi^+$	seen	—	—
$[\pi^+ \pi^- \pi^0]_D K^-$	( 4.6 $\pm$ 0.9 ) $\times 10^{-6}$	—	—
$[K_S^0 K^+ \pi^-]_D K^+$	seen	—	—
$[K_S^0 K^- \pi^+]_D K^+$	seen	—	—
$[K^*(892)^+ K^-]_D K^+$	seen	—	—
$[K_S^0 K^- \pi^+]_D \pi^+$	seen	—	—
$[K^*(892)^+ K^-]_D \pi^+$	seen	—	—
$[K_S^0 K^+ \pi^-]_D \pi^+$	seen	—	—
$[K^*(892)^+ K^+]_D \pi^+$	seen	—	—
$\bar{D}^0 K^*(892)^+$	( 5.3 $\pm$ 0.4 ) $\times 10^{-4}$	2213	—
$D_{CP(-1)} K^*(892)^+$	[uuu] ( 2.7 $\pm$ 0.8 ) $\times 10^{-4}$	—	—
$D_{CP(+1)} K^*(892)^+$	[uuu] ( 5.8 $\pm$ 1.1 ) $\times 10^{-4}$	—	—
$\bar{D}^0 K^+ \pi^+ \pi^-$	( 5.4 $\pm$ 2.2 ) $\times 10^{-4}$	2237	—
$\bar{D}^0 K^+ \bar{K}^0$	( 5.5 $\pm$ 1.6 ) $\times 10^{-4}$	2189	—
$\bar{D}^0 K^+ \bar{K}^*(892)^0$	( 7.5 $\pm$ 1.7 ) $\times 10^{-4}$	2072	—
$\bar{D}^0 \pi^+ \pi^+ \pi^-$	( 5.7 $\pm$ 2.2 ) $\times 10^{-3}$	S=3.6 2289	—
$\bar{D}^0 \pi^+ \pi^+ \pi^-$ nonresonant	( 5 $\pm$ 4 ) $\times 10^{-3}$	2289	—
$\bar{D}^0 \pi^+ \rho^0$	( 4.2 $\pm$ 3.0 ) $\times 10^{-3}$	2208	—
$\bar{D}^0 a_1(1260)^+$	( 4 $\pm$ 4 ) $\times 10^{-3}$	2123	—
$\bar{D}^0 \omega \pi^+$	( 4.1 $\pm$ 0.9 ) $\times 10^{-3}$	2206	—
$D^*(2010)^- \pi^+ \pi^+$	( 1.35 $\pm$ 0.22 ) $\times 10^{-3}$	2247	—
$\bar{D}_1(2420)^0 \pi^+, \bar{D}_1^0 \rightarrow$	( 5.3 $\pm$ 2.3 ) $\times 10^{-4}$	2081	—
$D^*(2010)^- \pi^+$	—	—	—
$D^- \pi^+ \pi^+$	( 1.07 $\pm$ 0.05 ) $\times 10^{-3}$	2299	—
$D^- K^+ \pi^+$	( 7.7 $\pm$ 0.5 ) $\times 10^{-5}$	2260	—
$D_0^*(2400)^0 K^+, D_0^{*0} \rightarrow$	( 6.1 $\pm$ 2.4 ) $\times 10^{-4}$	—	—
$D_1^{*-} \pi^+$	( 3.6 $\pm$ 1.2 ) $\times 10^{-4}$	—	—
$D_1^{*0}(2760)^0 K^+, D_1^{*0} \rightarrow$	( 3.6 $\pm$ 1.2 ) $\times 10^{-4}$	—	—
$D_2^{*-} \pi^+$	( 2.32 $\pm$ 0.23 ) $\times 10^{-3}$	—	—
$D_2^{*0}(2460)^0 K^+, D_2^{*0} \rightarrow$	—	—	—
$D^+ K^0$	< 2.9 $\times 10^{-6}$	CL=90% 2278	—
$D^+ K^{*0}$	< 1.8 $\times 10^{-6}$	CL=90% 2211	—
$D^+ \bar{K}^{*0}$	< 1.4 $\times 10^{-6}$	CL=90% 2211	—
$\bar{D}^*(2007)^0 \pi^+$	( 5.18 $\pm$ 0.26 ) $\times 10^{-3}$	2256	—
$\bar{D}_{CP(+1)}^{*0} \pi^+$	[xxx] ( 2.9 $\pm$ 0.7 ) $\times 10^{-3}$	—	—
$\bar{D}_{CP(-1)}^{*0} \pi^+$	[xxx] ( 2.6 $\pm$ 1.0 ) $\times 10^{-3}$	—	—
$\bar{D}^*(2007)^0 \omega \pi^+$	( 4.5 $\pm$ 1.2 ) $\times 10^{-3}$	2149	—
$\bar{D}^*(2007)^0 \rho^+$	( 9.8 $\pm$ 1.7 ) $\times 10^{-3}$	2181	—
$\bar{D}^*(2007)^0 K^+$	( 4.20 $\pm$ 0.34 ) $\times 10^{-4}$	2227	—
$\bar{D}_{CP(+1)}^{*0} K^+$	[xxx] ( 2.8 $\pm$ 0.4 ) $\times 10^{-4}$	—	—
$\bar{D}_{CP(-1)}^{*0} K^+$	[xxx] ( 2.31 $\pm$ 0.33 ) $\times 10^{-4}$	—	—
$\bar{D}^*(2007)^0 K^*(892)^+$	( 8.1 $\pm$ 1.4 ) $\times 10^{-4}$	2156	—
$\bar{D}^*(2007)^0 K^+ \bar{K}^0$	< 1.06 $\times 10^{-3}$	CL=90% 2132	—
$\bar{D}^*(2007)^0 K^+ K^*(892)^0$	( 1.5 $\pm$ 0.4 ) $\times 10^{-3}$	2009	—
$\bar{D}^*(2007)^0 \pi^+ \pi^+ \pi^-$	( 1.03 $\pm$ 0.12 ) %	2236	—
$\bar{D}^*(2007)^0 a_1(1260)^+$	( 1.9 $\pm$ 0.5 ) %	2063	—
$\bar{D}^*(2007)^0 \pi^- \pi^+ \pi^+ \pi^0$	( 1.8 $\pm$ 0.4 ) %	2219	—
$\bar{D}^{*0} 3\pi^+ 2\pi^-$	( 5.7 $\pm$ 1.2 ) $\times 10^{-3}$	2196	—
$D^*(2010)^+ \pi^0$	< 3.6 $\times 10^{-6}$	2255	—
$D^*(2010)^+ K^0$	< 9.0 $\times 10^{-6}$	CL=90% 2225	—
$D^*(2010)^- \pi^+ \pi^+ \pi^0$	( 1.5 $\pm$ 0.7 ) %	2235	—
$D^*(2010)^- \pi^+ \pi^+ \pi^+ \pi^-$	( 2.6 $\pm$ 0.4 ) $\times 10^{-3}$	2217	—
$\bar{D}^{*0} \pi^+$	[yyy] ( 5.9 $\pm$ 1.3 ) $\times 10^{-3}$	—	—
$\bar{D}_1^*(2420)^0 \pi^+$	( 1.5 $\pm$ 0.6 ) $\times 10^{-3}$	S=1.3 2082	—



## Meson Summary Table

$\overline{D}_1(2420)^0 \pi^+ \times \text{B}(\overline{D}_1^0 \rightarrow \overline{D}^0 \pi^+ \pi^-)$	$(2.5 \pm 1.6) \times 10^{-4}$	S=4.0	2082	$D_s^{(*)+} \overline{D}^{*0}$	$(2.7 \pm 1.2) \%$	–
$\overline{D}_1(2420)^0 \pi^+ \times \text{B}(\overline{D}_1^0 \rightarrow \overline{D}^0 \pi^+ \pi^- \text{ (nonresonant)})$	$(2.3 \pm 1.0) \times 10^{-4}$		2082	$\overline{D}^* (2007)^0 D^* (2010)^+$	$(8.1 \pm 1.7) \times 10^{-4}$	1713
$\overline{D}_2^*(2462)^0 \pi^+ \times \text{B}(\overline{D}_2^{*0} \rightarrow D^- \pi^+)$	$(3.5 \pm 0.4) \times 10^{-4}$		–	$\overline{D}^0 D^* (2010)^+ + \overline{D}^* (2007)^0 D^+$	$< 1.30 \%$	CL=90% 1792
$\overline{D}_2^*(2462)^0 \pi^+ \times \text{B}(\overline{D}_2^{*0} \rightarrow \overline{D}^0 \pi^- \pi^+)$	$(2.3 \pm 1.1) \times 10^{-4}$		–	$\overline{D}^0 D^* (2010)^+$	$(3.9 \pm 0.5) \times 10^{-4}$	1792
$\overline{D}_2^*(2462)^0 \pi^+ \times \text{B}(\overline{D}_2^{*0} \rightarrow \overline{D}^0 \pi^- \pi^+ \text{ (nonresonant)})$	$< 1.7 \times 10^{-4}$	CL=90%	–	$\overline{D}^0 D^+$	$(3.8 \pm 0.4) \times 10^{-4}$	1866
$\overline{D}_2^*(2462)^0 \pi^+ \times \text{B}(\overline{D}_2^{*0} \rightarrow D^* (2010)^- \pi^+)$	$(2.2 \pm 1.1) \times 10^{-4}$		–	$\overline{D}^0 D^+ K^0$	$(1.55 \pm 0.21) \times 10^{-3}$	1571
$\overline{D}_0^*(2400)^0 \pi^+ \times \text{B}(\overline{D}_0^{*0} \rightarrow D^- \pi^+)$	$(6.4 \pm 1.4) \times 10^{-4}$		2128	$D^+ \overline{D}^* (2007)^0$	$(6.3 \pm 1.7) \times 10^{-4}$	1791
$\overline{D}_1(2421)^0 \pi^+ \times \text{B}(\overline{D}_1(2421)^0 \rightarrow D^{*-} \pi^+)$	$(6.8 \pm 1.5) \times 10^{-4}$		–	$\overline{D}^* (2007)^0 D^+ K^0$	$(2.1 \pm 0.5) \times 10^{-3}$	1475
$\overline{D}_2^*(2462)^0 \pi^+ \times \text{B}(\overline{D}_2^*(2462)^0 \rightarrow D^{*-} \pi^+)$	$(1.8 \pm 0.5) \times 10^{-4}$		–	$\overline{D}^0 D^* (2010)^+ K^0$	$(3.8 \pm 0.4) \times 10^{-3}$	1476
$\overline{D}_1'(2427)^0 \pi^+ \times \text{B}(\overline{D}_1'(2427)^0 \rightarrow D^{*-} \pi^+)$	$(5.0 \pm 1.2) \times 10^{-4}$		–	$\overline{D}^* (2007)^0 D^* (2010)^+ K^0$	$(9.2 \pm 1.2) \times 10^{-3}$	1362
$\overline{D}_1(2420)^0 \pi^+ \times \text{B}(\overline{D}_1^0 \rightarrow \overline{D}^{*0} \pi^+ \pi^-)$	$< 6 \times 10^{-6}$	CL=90%	2082	$\overline{D}^0 D^0 K^+$	$(1.45 \pm 0.33) \times 10^{-3}$	S=2.6 1577
$\overline{D}_1^*(2420)^0 \rho^+$	$< 1.4 \times 10^{-3}$	CL=90%	1996	$\overline{D}^* (2007)^0 D^0 K^+$	$(2.26 \pm 0.23) \times 10^{-3}$	1481
$\overline{D}_2^*(2460)^0 \pi^+$	$< 1.3 \times 10^{-3}$	CL=90%	2063	$\overline{D}^0 D^* (2007)^0 K^+$	$(6.3 \pm 0.5) \times 10^{-3}$	1481
$\overline{D}_2^*(2460)^0 \pi^+ \times \text{B}(\overline{D}_2^{*0} \rightarrow \overline{D}^{*0} \pi^+ \pi^-)$	$< 2.2 \times 10^{-5}$	CL=90%	2063	$\overline{D}^* (2007)^0 D^* (2007)^0 K^+$	$(1.12 \pm 0.13) \%$	1368
$\overline{D}_2^*(2460)^0 \rho^+$	$< 4.7 \times 10^{-3}$	CL=90%	1977	$D^- D^+ K^+$	$(2.2 \pm 0.7) \times 10^{-4}$	1571
$\overline{D}_s^0 D_s^+$	$(9.0 \pm 0.9) \times 10^{-3}$		1815	$D^- D^* (2010)^+ K^+$	$(6.3 \pm 1.1) \times 10^{-4}$	1475
$D_{s0}^* (2317)^+ \overline{D}^0, D_{s0}^{*+} \rightarrow D_s^+ \pi^0$	$(7.9 \pm 1.5) \times 10^{-4}$		1605	$D^* (2010)^- D^+ K^+$	$(6.0 \pm 1.3) \times 10^{-4}$	1475
$D_{s0} (2317)^+ \overline{D}^0 \times \text{B}(D_{s0} (2317)^+ \rightarrow D_s^{*+} \gamma)$	$< 7.6 \times 10^{-4}$	CL=90%	1605	$D^* (2010)^- D^* (2010)^+ K^+$	$(1.32 \pm 0.18) \times 10^{-3}$	1363
$D_{s0} (2317)^+ \overline{D}^* (2007)^0 \times \text{B}(D_{s0} (2317)^+ \rightarrow D_s^+ \pi^0)$	$(9 \pm 7) \times 10^{-4}$		1511	$(\overline{D} + \overline{D}^*) (D + D^*) K$	$(4.05 \pm 0.30) \%$	–
$D_{sJ} (2457)^+ \overline{D}^0$	$(3.1 \pm 0.9) \times 10^{-3}$		–	$D_s^+ \pi^0$	$(1.6 \pm 0.5) \times 10^{-5}$	2270
$D_{sJ} (2457)^+ \overline{D}^0 \times \text{B}(D_{sJ} (2457)^+ \rightarrow D_s^+ \gamma)$	$(4.6 \pm 1.3) \times 10^{-4}$		–	$D_s^+ \pi^0$	$< 2.6 \times 10^{-4}$	CL=90% 2215
$D_{sJ} (2457)^+ \overline{D}^0 \times \text{B}(D_{sJ} (2457)^+ \rightarrow D_s^+ \pi^+ \pi^-)$	$< 2.2 \times 10^{-4}$	CL=90%	–	$D_s^+ \eta$	$< 4 \times 10^{-4}$	CL=90% 2235
$D_{sJ} (2457)^+ \overline{D}^0 \times \text{B}(D_{sJ} (2457)^+ \rightarrow D_s^+ \pi^+ \pi^-)$	$< 2.7 \times 10^{-4}$	CL=90%	–	$D_s^{*+} \eta$	$< 6 \times 10^{-4}$	CL=90% 2178
$D_{sJ} (2457)^+ \overline{D}^0 \times \text{B}(D_{sJ} (2457)^+ \rightarrow D_s^+ \pi^0)$	$< 9.8 \times 10^{-4}$	CL=90%	–	$D_s^{*+} \rho^0$	$< 3.0 \times 10^{-4}$	CL=90% 2197
$D_{sJ} (2457)^+ \overline{D}^* (2007)^0 \times \text{B}(D_{sJ} (2457)^+ \rightarrow D_s^+ \gamma)$	$(1.20 \pm 0.30) \%$		–	$D_s^{*+} \rho^0$	$< 4 \times 10^{-4}$	CL=90% 2138
$D_{sJ} (2457)^+ \overline{D}^* (2007)^0 \times \text{B}(D_{sJ} (2457)^+ \rightarrow D_s^+ \gamma)$	$(1.4 \pm 0.7) \times 10^{-3}$		–	$D_s^+ \omega$	$< 4 \times 10^{-4}$	CL=90% 2195
$\overline{D}^0 D_{s1} (2536)^+ \times \text{B}(D_{s1} (2536)^+ \rightarrow D^* (2007)^0 K^+ + D^* (2010)^+ K^0)$	$(4.0 \pm 1.0) \times 10^{-4}$		1447	$D_s^{*+} \omega$	$< 6 \times 10^{-4}$	CL=90% 2136
$\overline{D}^0 D_{s1} (2536)^+ \times \text{B}(D_{s1} (2536)^+ \rightarrow D^* (2007)^0 K^+ + D^* (2010)^+ K^0)$	$(2.2 \pm 0.7) \times 10^{-4}$		1447	$D_s^+ a_1(1260)^0$	$< 1.8 \times 10^{-3}$	CL=90% 2079
$\overline{D}^0 D_{s1} (2536)^+ \times \text{B}(D_{s1} (2536)^+ \rightarrow D^* (2007)^0 K^+)$	$(5.5 \pm 1.6) \times 10^{-4}$		1339	$D_s^{*+} a_1(1260)^0$	$< 1.3 \times 10^{-3}$	CL=90% 2015
$\overline{D}^0 D_{s1} (2536)^+ \times \text{B}(D_{s1} (2536)^+ \rightarrow D^* (2007)^0 K^+)$	$(2.3 \pm 1.1) \times 10^{-4}$		1447	$D_s^+ \phi$	$(1.7 \pm 1.2) \times 10^{-6}$	2141
$\overline{D}^0 D_{s1} (2536)^+ \times \text{B}(D_{sJ} (2700)^+ \rightarrow D^0 K^+)$	$(5.6 \pm 1.8) \times 10^{-4}$	S=1.7	–	$D_s^{*+} \phi$	$< 1.2 \times 10^{-5}$	CL=90% 2079
$\overline{D}^{*0} D_{s1} (2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0$	$(3.9 \pm 2.6) \times 10^{-4}$		1339	$D_s^+ \overline{K}^0$	$< 8 \times 10^{-4}$	CL=90% 2242
$\overline{D}^0 D_{sJ} (2573)^+, D_{sJ}^+ \rightarrow D^0 K^+$	$(8 \pm 15) \times 10^{-6}$		–	$D_s^{*+} \overline{K}^0$	$< 9 \times 10^{-4}$	CL=90% 2185
$\overline{D}^{*0} D_{sJ} (2573), D_{sJ}^+ \rightarrow D^0 K^+$	$< 2 \times 10^{-4}$	CL=90%	1306	$D_s^+ \overline{K}^* (892)^0$	$< 4.4 \times 10^{-6}$	CL=90% 2172
$\overline{D}^* (2007)^0 D_{sJ} (2573), D_{sJ}^+ \rightarrow D^0 K^+$	$< 5 \times 10^{-4}$	CL=90%	1306	$D_s^+ K^{*0}$	$< 3.5 \times 10^{-6}$	CL=90% 2172
$\overline{D}^0 D_s^{*+}$	$(7.6 \pm 1.6) \times 10^{-3}$		1734	$D_s^{*+} \overline{K}^* (892)^0$	$< 3.5 \times 10^{-4}$	CL=90% 2112
$\overline{D}^* (2007)^0 D_s^+$	$(8.2 \pm 1.7) \times 10^{-3}$		1737	$D_s^- \pi^+ K^+$	$(1.80 \pm 0.22) \times 10^{-4}$	2222
$\overline{D}^* (2007)^0 D_s^+$	$(1.71 \pm 0.24) \%$		1651	$D_s^{*-} \pi^+ K^+$	$(1.45 \pm 0.24) \times 10^{-4}$	2164
				$D_s^- \pi^+ K^* (892)^+$	$< 5 \times 10^{-3}$	CL=90% 2138
				$D_s^{*-} \pi^+ K^* (892)^+$	$< 7 \times 10^{-3}$	CL=90% 2076
				$D_s^- K^+ K^+$	$(9.7 \pm 2.1) \times 10^{-6}$	2149
				$D_s^{*-} K^+ K^+$	$< 1.5 \times 10^{-5}$	CL=90% 2088
				<b>Charmonium modes</b>		
				$\eta_c K^+$	$(9.6 \pm 1.1) \times 10^{-4}$	1751
				$\eta_c K^+, \eta_c \rightarrow K_S^0 K^\mp \pi^\pm$	$(2.7 \pm 0.6) \times 10^{-5}$	–
				$\eta_c K^* (892)^+$	$(1.0 \pm 0.5) \times 10^{-3}$	1646
				$\eta_c K^+ \pi^+ \pi^-$	$< 3.9 \times 10^{-4}$	CL=90% 1684
				$\eta_c K^+ \omega (782)$	$< 5.3 \times 10^{-4}$	CL=90% 1476
				$\eta_c K^+ \eta$	$< 2.2 \times 10^{-4}$	CL=90% 1588
				$\eta_c K^+ \pi^0$	$< 6.2 \times 10^{-5}$	CL=90% 1723
				$\eta_c (2S) K^+$	$(3.4 \pm 1.8) \times 10^{-4}$	1319
				$\eta_c (2S) K^+, \eta_c \rightarrow p \overline{p}$	$< 1.06 \times 10^{-7}$	CL=95% –
				$\eta_c (2S) K^+, \eta_c \rightarrow K_S^0 K^\mp \pi^\pm$	$(3.4 \pm 2.3) \times 10^{-6}$	–
				$h_c(1P) K^+, h_c \rightarrow J/\psi \pi^+ \pi^-$	$< 3.4 \times 10^{-6}$	CL=90% 1401
				$X(3730)^0 K^+, X^0 \rightarrow \eta_c \eta$	$< 4.6 \times 10^{-5}$	CL=90% –
				$X(3730)^0 K^+, X^0 \rightarrow \eta_c \pi^0$	$< 5.7 \times 10^{-6}$	CL=90% –
				$X(3872) K^+$	$< 3.2 \times 10^{-4}$	CL=90% 1141
				$X(3872) K^+, X \rightarrow p \overline{p}$	$< 1.7 \times 10^{-8}$	CL=95% –
				$X(3872) K^+, X \rightarrow J/\psi \pi^+ \pi^-$	$(8.6 \pm 0.8) \times 10^{-6}$	1141
				$X(3872) K^+, X \rightarrow J/\psi \gamma$	$(2.1 \pm 0.4) \times 10^{-6}$	S=1.1 1141
				$X(3872) K^+, X \rightarrow \psi(2S) \gamma$	$(4 \pm 4) \times 10^{-6}$	S=2.5 1141
				$X(3872) K^+, X \rightarrow J/\psi (1S) \eta$	$< 7.7 \times 10^{-6}$	CL=90% 1141
				$X(3872) K^+, X \rightarrow D^0 \overline{D}^0$	$< 6.0 \times 10^{-5}$	CL=90% 1141
				$X(3872) K^+, X \rightarrow D^+ D^-$	$< 4.0 \times 10^{-5}$	CL=90% 1141
				$X(3872) K^+, X \rightarrow D^0 \overline{D}^0 \pi^0$	$(1.0 \pm 0.4) \times 10^{-4}$	1141
				$X(3872) K^+, X \rightarrow \overline{D}^{*0} D^0$	$(8.5 \pm 2.6) \times 10^{-5}$	S=1.4 1141
				$X(3872)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-$	$< 3.0 \times 10^{-5}$	CL=90% –

## Meson Summary Table

					<b>K or K* modes</b>		
$X(3872)^0 K^+, X^0 \rightarrow$	$< 6.9$	$\times 10^{-5}$	CL=90%	—	$K^0 \pi^+$	$(2.37 \pm 0.08) \times 10^{-5}$	2614
$\eta_c \omega(782)$					$K^+ \pi^0$	$(1.29 \pm 0.05) \times 10^{-5}$	2615
$X(3915)^0 K^+, X^0 \rightarrow \eta_c \eta$	$< 3.3$	$\times 10^{-5}$	CL=90%	—	$\eta' K^+$	$(7.06 \pm 0.25) \times 10^{-5}$	2528
$X(3915)^0 K^+, X^0 \rightarrow \eta_c \pi^0$	$< 1.8$	$\times 10^{-5}$	CL=90%	—	$\eta' K^*(892)^+$	$(4.8 \pm 1.8) \times 10^{-6}$	2472
$X(4014)^0 K^+, X^0 \rightarrow \eta_c \eta$	$< 3.9$	$\times 10^{-5}$	CL=90%	—	$\eta' K_0^*(1430)^+$	$(5.2 \pm 2.1) \times 10^{-6}$	—
$X(4014)^0 K^+, X^0 \rightarrow \eta_c \pi^0$	$< 1.2$	$\times 10^{-5}$	CL=90%	—	$\eta' K_2^*(1430)^+$	$(2.8 \pm 0.5) \times 10^{-5}$	2346
$X(3900)^0 K^+, X^0 \rightarrow$	$< 4.7$	$\times 10^{-5}$	CL=90%	—	$\eta K^+$	$(2.4 \pm 0.4) \times 10^{-6}$	S=1.7 2588
$\eta_c \pi^+ \pi^-$					$\eta K^*(892)^+$	$(1.93 \pm 0.16) \times 10^{-5}$	2534
$X(4020)^0 K^+, X^0 \rightarrow$	$< 1.6$	$\times 10^{-5}$	CL=90%	—	$\eta K_0^*(1430)^+$	$(1.8 \pm 0.4) \times 10^{-5}$	—
$\eta_c \pi^+ \pi^-$					$\eta K_2^*(1430)^+$	$(9.1 \pm 3.0) \times 10^{-6}$	2414
$X(3872) K^*(892)^+, X \rightarrow$	$< 4.8$	$\times 10^{-6}$	CL=90%	939	$\eta(1295) K^+ \times B(\eta(1295) \rightarrow$	$(2.9 \pm 0.8) \times 10^{-6}$	2455
$J/\psi \gamma$					$\eta \pi \pi$		
$X(3872) K^*(892)^+, X \rightarrow$	$< 2.8$	$\times 10^{-5}$	CL=90%	939	$\eta(1405) K^+ \times B(\eta(1405) \rightarrow$	$< 1.3$	$\times 10^{-6}$ CL=90% 2425
$\psi(2S) \gamma$					$\eta \pi \pi$		
$X(3872)^+ K^0, X^+ \rightarrow$	[zzz] $< 6.1$	$\times 10^{-6}$	CL=90%	—	$\eta(1405) K^+ \times B(\eta(1405) \rightarrow$	$< 1.2$	$\times 10^{-6}$ CL=90% 2425
$J/\psi(1S) \pi^+ \pi^0$					$K^* K$		
$X(3872) K^0 \pi^+, X \rightarrow$	$(1.06 \pm 0.31) \times 10^{-5}$			—	$\eta(1475) K^+ \times B(\eta(1475) \rightarrow$	$(1.38 \pm 0.21) \times 10^{-5}$	2406
$J/\psi(1S) \pi^+ \pi^-$					$K^* K$		
$X(4430)^+ K^0, X^+ \rightarrow J/\psi \pi^+$	$< 1.5$	$\times 10^{-5}$	CL=95%	—	$f_1(1285) K^+$	$< 2.0$	$\times 10^{-6}$ CL=90% 2458
$X(4430)^+ K^0, X^+ \rightarrow$	$< 4.7$	$\times 10^{-5}$	CL=95%	—	$f_1(1420) K^+ \times B(f_1(1420) \rightarrow$	$< 2.9$	$\times 10^{-6}$ CL=90% 2420
$\psi(2S) \pi^+$					$\eta \pi \pi$		
$X(4260)^0 K^+, X^0 \rightarrow$	$< 2.9$	$\times 10^{-5}$	CL=95%	—	$f_1(1420) K^+ \times B(f_1(1420) \rightarrow$	$< 4.1$	$\times 10^{-6}$ CL=90% 2420
$J/\psi \pi^+ \pi^-$					$K^* K$		
$X(3915) K^+, X \rightarrow J/\psi \gamma$	$< 1.4$	$\times 10^{-5}$	CL=90%	—	$\phi(1680) K^+ \times B(\phi(1680) \rightarrow$	$< 3.4$	$\times 10^{-6}$ CL=90% 2344
$X(3930)^0 K^+, X^0 \rightarrow J/\psi \gamma$	$< 2.5$	$\times 10^{-6}$	CL=90%	—	$K^* K$		
$J/\psi(1S) K^+$	$(1.026 \pm 0.031) \times 10^{-3}$			1684	$f_0(1500) K^+$	$(3.7 \pm 2.2) \times 10^{-6}$	2398
$J/\psi(1S) K^+ \pi^+ \pi^-$	$(8.1 \pm 1.3) \times 10^{-4}$		S=2.5	1612	$\omega K^+$	$(6.5 \pm 0.4) \times 10^{-6}$	2558
$J/\psi(1S) K^+ K^- K^+$	$(3.37 \pm 0.29) \times 10^{-5}$			1252	$\omega K^*(892)^+$	$< 7.4$	$\times 10^{-6}$ CL=90% 2503
$X(3915) K^+, X \rightarrow p \bar{p}$	$< 7.1$	$\times 10^{-8}$	CL=95%	—	$\omega(K\pi)_0^{*+}$	$(2.8 \pm 0.4) \times 10^{-5}$	—
$J/\psi(1S) K^*(892)^+$	$(1.43 \pm 0.08) \times 10^{-3}$			1571	$\omega K_0^*(1430)^+$	$(2.4 \pm 0.5) \times 10^{-5}$	—
$J/\psi(1S) K(1270)^+$	$(1.8 \pm 0.5) \times 10^{-3}$			1390	$\omega K_2^*(1430)^+$	$(2.1 \pm 0.4) \times 10^{-5}$	2380
$J/\psi(1S) K(1400)^+$	$< 5$	$\times 10^{-4}$	CL=90%	1308	$a_0(980)^+ K^0 \times B(a_0(980)^+ \rightarrow$	$< 3.9$	$\times 10^{-6}$ CL=90% —
$J/\psi(1S) \eta K^+$	$(1.24 \pm 0.14) \times 10^{-4}$			1510	$\eta \pi^+$		
$X^{c-odd}(3872) K^+,$	$< 3.8$	$\times 10^{-6}$	CL=90%	—	$a_0(980)^0 K^+ \times B(a_0(980)^0 \rightarrow$	$< 2.5$	$\times 10^{-6}$ CL=90% —
$X^{c-odd} \rightarrow J/\psi \eta$					$\eta \pi^0$		
$\psi(4160) K^+, \psi \rightarrow J/\psi \eta$	$< 7.4$	$\times 10^{-6}$	CL=90%	—	$K^*(892)^0 \pi^+$	$(1.01 \pm 0.09) \times 10^{-5}$	2562
$J/\psi(1S) \eta' K^+$	$< 8.8$	$\times 10^{-5}$	CL=90%	1273	$K^*(892)^+ \pi^0$	$(8.2 \pm 1.9) \times 10^{-6}$	2563
$J/\psi(1S) \phi K^+$	$(5.0 \pm 0.4) \times 10^{-5}$			1227	$K^+ \pi^- \pi^+$	$(5.10 \pm 0.29) \times 10^{-5}$	2609
$X(4140) K^+, X \rightarrow$	$(10 \pm 4) \times 10^{-6}$			—	$K^+ \pi^- \pi^+ \text{ nonresonant}$	$(1.63 \pm 0.21) \times 10^{-5}$	2609
$J/\psi(1S) \phi$					$\omega(782) K^+$	$(6 \pm 9) \times 10^{-6}$	2558
$X(4274) K^+, X \rightarrow$	$< 4$	$\times 10^{-6}$	CL=90%	—	$K^+ f_0(980) \times B(f_0(980) \rightarrow$	$(9.4 \pm 1.0) \times 10^{-6}$	2522
$J/\psi(1S) \phi$					$\pi^+ \pi^-$		
$J/\psi(1S) \omega K^+$	$(3.20 \pm 0.60) \times 10^{-4}$			1388	$f_2(1270)^0 K^+$	$(1.07 \pm 0.27) \times 10^{-6}$	—
$X(3872) K^+, X \rightarrow J/\psi \omega$	$(6.0 \pm 2.2) \times 10^{-6}$			1141	$f_0(1370)^0 K^+ \times$	$< 1.07$	$\times 10^{-5}$ CL=90% —
$X(3915) K^+, X \rightarrow J/\psi \omega$	$(3.0 \pm 0.9) \times 10^{-5}$			1103	$B(f_0(1370)^0 \rightarrow \pi^+ \pi^-)$		
$J/\psi(1S) \pi^+$	$(4.1 \pm 0.4) \times 10^{-5}$		S=2.6	1728	$\rho^0(1450) K^+ \times$	$< 1.17$	$\times 10^{-5}$ CL=90% —
$J/\psi(1S) \rho^+$	$(5.0 \pm 0.8) \times 10^{-5}$			1611	$B(\rho^0(1450) \rightarrow \pi^+ \pi^-)$		
$J/\psi(1S) \pi^+ \pi^0 \text{ nonresonant}$	$< 7.3$	$\times 10^{-6}$	CL=90%	1717	$f_2'(1525) K^+ \times$	$< 3.4$	$\times 10^{-6}$ CL=90% 2392
$J/\psi(1S) a_1(1260)^+$	$< 1.2$	$\times 10^{-3}$	CL=90%	1415	$B(f_2'(1525) \rightarrow \pi^+ \pi^-)$		
$J/\psi p \bar{p} \pi^+$	$< 5.0$	$\times 10^{-7}$	CL=90%	643	$K^+ \rho^0$	$(3.7 \pm 0.5) \times 10^{-6}$	2559
$J/\psi(1S) p \bar{\Lambda}$	$(1.18 \pm 0.31) \times 10^{-5}$			567	$K_0^*(1430)^0 \pi^+$	$(4.5 \pm 0.9) \times 10^{-5}$	S=1.5 2445
$J/\psi(1S) \bar{\Sigma}^0 p$	$< 1.1$	$\times 10^{-5}$	CL=90%	—	$K_2^*(1430)^0 \pi^+$	$(5.6 \pm 2.2) \times 10^{-6}$	2445
$J/\psi(1S) D^+$	$< 1.2$	$\times 10^{-4}$	CL=90%	871	$K^*(1410)^0 \pi^+$	$< 4.5$	$\times 10^{-5}$ CL=90% 2448
$J/\psi(1S) \bar{D}^0 \pi^+$	$< 2.5$	$\times 10^{-5}$	CL=90%	665	$K^*(1680)^0 \pi^+$	$< 1.2$	$\times 10^{-5}$ CL=90% 2358
$\psi(2S) \pi^+$	$(2.44 \pm 0.30) \times 10^{-5}$			1347	$K^+ \pi^0 \pi^0$	$(1.62 \pm 0.19) \times 10^{-5}$	2610
$\psi(2S) K^+$	$(6.26 \pm 0.24) \times 10^{-4}$			1284	$f_0(980) K^+ \times B(f_0 \rightarrow \pi^0 \pi^0)$	$(2.8 \pm 0.8) \times 10^{-6}$	2522
$\psi(2S) K^*(892)^+$	$(6.7 \pm 1.4) \times 10^{-4}$		S=1.3	1116	$K^- \pi^+ \pi^+$	$< 9.5$	$\times 10^{-7}$ CL=90% 2609
$\psi(2S) K^+ \pi^+ \pi^-$	$(4.3 \pm 0.5) \times 10^{-4}$			1179	$K^- \pi^+ \pi^+ \text{ nonresonant}$	$< 5.6$	$\times 10^{-5}$ CL=90% 2609
$\psi(3770) K^+$	$(4.9 \pm 1.3) \times 10^{-4}$			1218	$K_1(1270)^0 \pi^+$	$< 4.0$	$\times 10^{-5}$ CL=90% 2484
$\psi(3770) K^+, \psi \rightarrow D^0 \bar{D}^0$	$(1.5 \pm 0.5) \times 10^{-4}$		S=1.4	1218	$K_1(1400)^0 \pi^+$	$< 3.9$	$\times 10^{-5}$ CL=90% 2451
$\psi(3770) K^+, \psi \rightarrow D^+ D^-$	$(9.4 \pm 3.5) \times 10^{-5}$			1218	$K^0 \pi^+ \pi^0$	$< 6.6$	$\times 10^{-5}$ CL=90% 2609
$\psi(4040) K^+$	$< 1.3$	$\times 10^{-4}$	CL=90%	1003	$K^0 \rho^+$	$(8.0 \pm 1.5) \times 10^{-6}$	2558
$\psi(4160) K^+$	$(5.1 \pm 2.7) \times 10^{-4}$			868	$K^*(892)^+ \pi^+ \pi^-$	$(7.5 \pm 1.0) \times 10^{-5}$	2557
$\psi(4160) K^+, \psi \rightarrow \bar{D}^0 D^0$	$(8 \pm 5) \times 10^{-5}$			—	$K^*(892)^+ \rho^0$	$(4.6 \pm 1.1) \times 10^{-6}$	2504
$X_{c0} \pi^+, X_{c0} \rightarrow \pi^+ \pi^-$	$< 1$	$\times 10^{-7}$	CL=90%	1531	$K^*(892)^+ f_0(980)$	$(4.2 \pm 0.7) \times 10^{-6}$	2466
$X_{c0}(1P) K^+$	$(1.50 \pm 0.15) \times 10^{-4}$			1478	$a_1^+ K^0$	$(3.5 \pm 0.7) \times 10^{-5}$	—
$X_{c0} K^*(892)^+$	$< 2.1$	$\times 10^{-4}$	CL=90%	1341	$b_1^+ K^0 \times B(b_1^+ \rightarrow \omega \pi^+)$	$(9.6 \pm 1.9) \times 10^{-6}$	—
$X_{c2} \pi^+, X_{c2} \rightarrow \pi^+ \pi^-$	$< 1$	$\times 10^{-7}$	CL=90%	1437	$K^*(892)^0 \rho^+$	$(9.2 \pm 1.5) \times 10^{-6}$	2504
$X_{c2} K^+$	$(1.1 \pm 0.4) \times 10^{-5}$			1379	$K_1(1400)^+ \rho^0$	$< 7.8$	$\times 10^{-4}$ CL=90% 2388
$X_{c2} K^*(892)^+$	$< 1.2$	$\times 10^{-4}$	CL=90%	1228	$K_2^*(1430)^+ \rho^0$	$< 1.5$	$\times 10^{-3}$ CL=90% 2381
$X_{c1}(1P) \pi^+$	$(2.2 \pm 0.5) \times 10^{-5}$			1468	$b_1^0 K^+ \times B(b_1^0 \rightarrow \omega \pi^0)$	$(9.1 \pm 2.0) \times 10^{-6}$	—
$X_{c1}(1P) K^+$	$(4.79 \pm 0.23) \times 10^{-4}$			1412	$b_1^+ K^* \times B(b_1^+ \rightarrow \omega \pi^+)$	$< 5.9$	$\times 10^{-6}$ CL=90% —
$X_{c1}(1P) K^*(892)^+$	$(3.0 \pm 0.6) \times 10^{-4}$		S=1.1	1265			
$h_c(1P) K^+$	$< 3.8$	$\times 10^{-5}$	CL=90%	1401			
$h_c(1P) K^+, h_c \rightarrow p \bar{p}$	$< 6.4$	$\times 10^{-8}$	CL=95%	—			

## Meson Summary Table

$b_1^0 K^{*+} \times B(b_1^0 \rightarrow \omega \pi^0)$	< 6.7	$\times 10^{-6}$	CL=90%	—	$\pi^+ \pi^- \pi^+$ nonresonant	( 5.3 $\pm$ 1.5 ) $\times 10^{-6}$	2630
$K_1^+ \bar{K}^0$	( 1.31 $\pm$ 0.17 )	$\times 10^{-6}$	S=1.2	2593	$\pi^+ \pi^0 \pi^0$	< 8.9	$\times 10^{-4}$ CL=90% 2631
$\bar{K}^0 K^+ \pi^0$	< 2.4	$\times 10^{-5}$	CL=90%	2578	$\rho^+ \pi^0$	( 1.09 $\pm$ 0.14 )	$\times 10^{-5}$ 2581
$K^+ K_S^0 K_S^0$	( 1.08 $\pm$ 0.06 )	$\times 10^{-5}$		2521	$\pi^+ \pi^- \pi^+ \pi^0$	< 4.0	$\times 10^{-3}$ CL=90% 2622
$f_0(980) K^+, f_0 \rightarrow K_S^0 K_S^0$	( 1.47 $\pm$ 0.33 )	$\times 10^{-5}$		—	$\rho^+ \rho^0$	( 2.40 $\pm$ 0.19 )	$\times 10^{-5}$ 2523
$f_0(1710) K^+, f_0 \rightarrow K_S^0 K_S^0$	( 4.8 $\pm$ 4.0 )	$\times 10^{-7}$		—	$\rho^+ f_0(980), f_0 \rightarrow \pi^+ \pi^-$	< 2.0	$\times 10^{-6}$ CL=90% 2486
$K^+ K_S^0 K_S^0$ nonresonant	( 2.0 $\pm$ 0.4 )	$\times 10^{-5}$		2521	$a_1(1260)^+ \pi^0$	( 2.6 $\pm$ 0.7 )	$\times 10^{-5}$ 2494
$K_S^0 K_S^0 \pi^+$	< 5.1	$\times 10^{-7}$	CL=90%	2577	$a_1(1260)^0 \pi^+$	( 2.0 $\pm$ 0.6 )	$\times 10^{-5}$ 2494
$K^+ K^- \pi^+$	( 5.0 $\pm$ 0.7 )	$\times 10^{-6}$		2578	$\omega \pi^+$	( 6.9 $\pm$ 0.5 )	$\times 10^{-6}$ 2580
$K^+ K^- \pi^+$ nonresonant	< 7.5	$\times 10^{-5}$	CL=90%	2578	$\omega \rho^+$	( 1.59 $\pm$ 0.21 )	$\times 10^{-5}$ 2522
$K^+ \bar{K}^*(892)^0$	< 1.1	$\times 10^{-6}$	CL=90%	2540	$\eta \pi^+$	( 4.02 $\pm$ 0.27 )	$\times 10^{-6}$ 2609
$K^+ \bar{K}_0^*(1430)^0$	< 2.2	$\times 10^{-6}$	CL=90%	2421	$\eta \rho^+$	( 7.0 $\pm$ 2.9 )	$\times 10^{-6}$ S=2.8 2553
$K^+ K^+ \pi^-$	< 1.6	$\times 10^{-7}$	CL=90%	2578	$\eta' \pi^+$	( 2.7 $\pm$ 0.9 )	$\times 10^{-6}$ S=1.9 2551
$K^+ K^+ \pi^-$ nonresonant	< 8.79	$\times 10^{-5}$	CL=90%	2578	$\eta' \rho^+$	( 9.7 $\pm$ 2.2 )	$\times 10^{-6}$ 2492
$f_2'(1525) K^+$	( 1.8 $\pm$ 0.5 )	$\times 10^{-6}$	S=1.1	2392	$\phi \pi^+$	< 1.5	$\times 10^{-7}$ CL=90% 2539
$K^{*+} \pi^+ K^-$	< 1.18	$\times 10^{-5}$	CL=90%	2524	$\phi \rho^+$	< 3.0	$\times 10^{-6}$ CL=90% 2480
$K^*(892)^+ K^*(892)^0$	( 9.1 $\pm$ 2.9 )	$\times 10^{-7}$		2484	$a_0(980)^0 \pi^+, a_0^0 \rightarrow \eta \pi^0$	< 5.8	$\times 10^{-6}$ CL=90% —
$K^{*+} K^+ \pi^-$	< 6.1	$\times 10^{-6}$	CL=90%	2524	$a_0(980)^+ \pi^0, a_0^+ \rightarrow \eta \pi^+$	< 1.4	$\times 10^{-6}$ CL=90% —
$K^+ K^- K^+$	( 3.40 $\pm$ 0.14 )	$\times 10^{-5}$	S=1.4	2523	$\pi^+ \pi^+ \pi^+ \pi^- \pi^-$	< 8.6	$\times 10^{-4}$ CL=90% 2608
$K^+ \phi$	( 8.8 $\pm$ 0.7 )	$\times 10^{-6}$	S=1.1	2516	$\rho^0 a_1(1260)^+$	< 6.2	$\times 10^{-4}$ CL=90% 2433
$f_0(980) K^+ \times B(f_0(980) \rightarrow K^+ K^-)$	( 9.4 $\pm$ 3.2 )	$\times 10^{-6}$		2522	$\rho^0 a_2(1320)^+$	< 7.2	$\times 10^{-4}$ CL=90% 2410
$a_2(1320) K^+ \times B(a_2(1320) \rightarrow K^+ K^-)$	< 1.1	$\times 10^{-6}$	CL=90%	2449	$b_0^0 \pi^+, b_1^0 \rightarrow \omega \pi^0$	( 6.7 $\pm$ 2.0 )	$\times 10^{-6}$ —
$X_0(1550) K^+ \times B(X_0(1550) \rightarrow K^+ K^-)$	( 4.3 $\pm$ 0.7 )	$\times 10^{-6}$		—	$b_1^+ \pi^0, b_1^+ \rightarrow \omega \pi^+$	< 3.3	$\times 10^{-6}$ CL=90% —
$\phi(1680) K^+ \times B(\phi(1680) \rightarrow K^+ K^-)$	< 8	$\times 10^{-7}$	CL=90%	2344	$\pi^+ \pi^+ \pi^+ \pi^- \pi^- \pi^0$	< 6.3	$\times 10^{-3}$ CL=90% 2592
$f_0(1710) K^+ \times B(f_0(1710) \rightarrow K^+ K^-)$	( 1.1 $\pm$ 0.6 )	$\times 10^{-6}$		2330	$b_1^+ \rho^0, b_1^+ \rightarrow \omega \pi^+$	< 5.2	$\times 10^{-6}$ CL=90% —
$K^+ K^- K^+$ nonresonant	( 2.38 $\pm$ 0.28 )	$\times 10^{-5}$		2523	$a_1(1260)^+ a_1(1260)^0$	< 1.3	% CL=90% 2336
$K^*(892)^+ K^+ K^-$	( 3.6 $\pm$ 0.5 )	$\times 10^{-5}$		2466	$b_1^0 \rho^+, b_1^0 \rightarrow \omega \pi^0$	< 3.3	$\times 10^{-6}$ CL=90% —
$K^*(892)^+ \phi$	( 10.0 $\pm$ 2.0 )	$\times 10^{-6}$	S=1.7	2460			
$\phi(K\pi)_0^{*+}$	( 8.3 $\pm$ 1.6 )	$\times 10^{-6}$		—			
$\phi K_1(1270)^+$	( 6.1 $\pm$ 1.9 )	$\times 10^{-6}$		2375			
$\phi K_1(1400)^+$	< 3.2	$\times 10^{-6}$	CL=90%	2339			
$\phi K^*(1410)^+$	< 4.3	$\times 10^{-6}$	CL=90%	—			
$\phi K_0^*(1430)^+$	( 7.0 $\pm$ 1.6 )	$\times 10^{-6}$		—			
$\phi K_2^*(1430)^+$	( 8.4 $\pm$ 2.1 )	$\times 10^{-6}$		2333			
$\phi K_2^*(1770)^+$	< 1.50	$\times 10^{-5}$	CL=90%	—			
$\phi K_2^*(1820)^+$	< 1.63	$\times 10^{-5}$	CL=90%	—			
$a_1^+ K^*0$	< 3.6	$\times 10^{-6}$	CL=90%	—			
$K^+ \phi \phi$	( 5.0 $\pm$ 1.2 )	$\times 10^{-6}$	S=2.3	2306			
$\eta' \eta' K^+$	< 2.5	$\times 10^{-5}$	CL=90%	2338			
$\omega \phi K^+$	< 1.9	$\times 10^{-6}$	CL=90%	2374			
$X(1812) K^+ \times B(X \rightarrow \omega \phi)$	< 3.2	$\times 10^{-7}$	CL=90%	—			
$K^*(892)^+ \gamma$	( 4.21 $\pm$ 0.18 )	$\times 10^{-5}$		2564			
$K_1(1270)^+ \gamma$	( 4.3 $\pm$ 1.3 )	$\times 10^{-5}$		2486			
$\eta K^+ \gamma$	( 7.9 $\pm$ 0.9 )	$\times 10^{-6}$		2588			
$\eta' K^+ \gamma$	( 2.9 $\pm$ 1.0 )	$\times 10^{-6}$		2528			
$\phi K^+ \gamma$	( 2.7 $\pm$ 0.4 )	$\times 10^{-6}$	S=1.2	2516			
$K^+ \pi^- \pi^+ \gamma$	( 2.76 $\pm$ 0.22 )	$\times 10^{-5}$	S=1.2	2609			
$K^*(892)^0 \pi^+ \gamma$	( 2.0 $\pm$ 0.7 )	$\times 10^{-5}$		2562			
$K^+ \rho^0 \gamma$	< 2.0	$\times 10^{-5}$	CL=90%	2559			
$K^+ \pi^- \pi^+ \gamma$ nonresonant	< 9.2	$\times 10^{-6}$	CL=90%	2609			
$K^0 \pi^+ \pi^0 \gamma$	( 4.6 $\pm$ 0.5 )	$\times 10^{-5}$		2609			
$K_1(1400)^+ \gamma$	< 1.5	$\times 10^{-5}$	CL=90%	2453			
$K_2^*(1430)^+ \gamma$	( 1.4 $\pm$ 0.4 )	$\times 10^{-5}$		2447			
$K^*(1680)^+ \gamma$	< 1.9	$\times 10^{-3}$	CL=90%	2360			
$K_3^*(1780)^+ \gamma$	< 3.9	$\times 10^{-5}$	CL=90%	2341			
$K_4^*(2045)^+ \gamma$	< 9.9	$\times 10^{-3}$	CL=90%	2244			
<b>Light unflavored meson modes</b>					<b>Baryon modes</b>		
$\rho^+ \gamma$	( 9.8 $\pm$ 2.5 )	$\times 10^{-7}$		2583	$p \bar{p} \pi^+$	( 1.62 $\pm$ 0.20 )	$\times 10^{-6}$ 2439
$\pi^+ \pi^0$	( 5.5 $\pm$ 0.4 )	$\times 10^{-6}$	S=1.2	2636	$p \bar{p} \pi^+$ nonresonant	< 5.3	$\times 10^{-5}$ CL=90% 2439
$\pi^+ \pi^+ \pi^-$	( 1.52 $\pm$ 0.14 )	$\times 10^{-5}$		2630	$p \bar{p} K^+$	( 5.9 $\pm$ 0.5 )	$\times 10^{-6}$ S=1.5 2348
$\rho^0 \pi^+$	( 8.3 $\pm$ 1.2 )	$\times 10^{-6}$		2581	$\Theta(1710)^{++} \bar{p}, \Theta^{++} \rightarrow p K^+$	[ <i>aaaa</i> ] < 9.1	$\times 10^{-8}$ CL=90% —
$\pi^+ f_0(980), f_0 \rightarrow \pi^+ \pi^-$	< 1.5	$\times 10^{-6}$	CL=90%	2545	$f_J(2220) K^+, f_J \rightarrow p \bar{p}$	[ <i>aaaa</i> ] < 4.1	$\times 10^{-7}$ CL=90% 2135
$\pi^+ f_2(1270)$	( 1.6 $\pm$ 0.7 )	$\times 10^{-6}$		2484	$p \bar{\Lambda}(1520)$	( 3.1 $\pm$ 0.6 )	$\times 10^{-7}$ 2322
$\rho(1450)^0 \pi^+, \rho^0 \rightarrow \pi^+ \pi^-$	( 1.4 $\pm$ 0.6 )	$\times 10^{-6}$		2434	$p \bar{p} K^+$ nonresonant	< 8.9	$\times 10^{-5}$ CL=90% 2348
$f_0(1370) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	< 4.0	$\times 10^{-6}$	CL=90%	2460	$p \bar{p} K^*(892)^+$	( 3.6 $\pm$ 0.8 )	$\times 10^{-6}$ 2215
$f_0(500) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	< 4.1	$\times 10^{-6}$	CL=90%	—	$f_J(2220) K^{*+}, f_J \rightarrow p \bar{p}$	< 7.7	$\times 10^{-7}$ CL=90% 2059
					$p \bar{\Lambda}$	< 3.2	$\times 10^{-7}$ CL=90% 2430
					$p \bar{\Lambda} \gamma$	( 2.4 $\pm$ 0.5 )	$\times 10^{-6}$ 2430
					$p \bar{\Lambda} \pi^0$	( 3.0 $\pm$ 0.7 )	$\times 10^{-6}$ 2402
					$p \bar{\Sigma}(1385)^0$	< 4.7	$\times 10^{-7}$ CL=90% 2362
					$\Delta^+ \bar{\Lambda}$	< 8.2	$\times 10^{-7}$ CL=90% —
					$p \bar{\Sigma} \gamma$	< 4.6	$\times 10^{-6}$ CL=90% 2413
					$p \bar{\Lambda} \pi^+ \pi^-$	( 5.9 $\pm$ 1.1 )	$\times 10^{-6}$ 2367
					$p \bar{\Lambda} \rho^0$	( 4.8 $\pm$ 0.9 )	$\times 10^{-6}$ 2214
					$p \bar{\Lambda} f_2(1270)$	( 2.0 $\pm$ 0.8 )	$\times 10^{-6}$ 2026
					$\Lambda \bar{\Lambda} \pi^+$	< 9.4	$\times 10^{-7}$ CL=90% 2358
					$\Lambda \bar{\Lambda} K^+$	( 3.4 $\pm$ 0.6 )	$\times 10^{-6}$ 2251
					$\Lambda \bar{\Lambda} K^{*+}$	( 2.2 $\pm$ 1.2 )	$\times 10^{-6}$ 2098
					$\bar{\Delta}^0 p$	< 1.38	$\times 10^{-6}$ CL=90% 2403
					$\Delta^{++} \bar{p}$	< 1.4	$\times 10^{-7}$ CL=90% 2403
					$D^+ p \bar{p}$	< 1.5	$\times 10^{-5}$ CL=90% 1860
					$D^*(2010)^+ p \bar{p}$	< 1.5	$\times 10^{-5}$ CL=90% 1786
					$\bar{D}^0 p \bar{p} \pi^+$	( 3.72 $\pm$ 0.27 )	$\times 10^{-4}$ 1789
					$\bar{D}^{*0} p \bar{p} \pi^+$	( 3.73 $\pm$ 0.32 )	$\times 10^{-4}$ 1709
					$D^- p \bar{p} \pi^+ \pi^-$	( 1.66 $\pm$ 0.30 )	$\times 10^{-4}$ 1705
					$D^{*-} p \bar{p} \pi^+ \pi^-$	( 1.86 $\pm$ 0.25 )	$\times 10^{-4}$ 1621
					$p \bar{\Lambda}^0 \bar{D}^0$	( 1.43 $\pm$ 0.32 )	$\times 10^{-5}$ —
					$p \bar{\Lambda}^0 \bar{D}^*(2007)^0$	< 5	$\times 10^{-5}$ CL=90% —
					$\bar{\Lambda}_c^- p \pi^+$	( 2.2 $\pm$ 0.4 )	$\times 10^{-4}$ S=2.2 1980
					$\bar{\Lambda}_c^- \Delta(1232)^{++}$	< 1.9	$\times 10^{-5}$ CL=90% 1928
					$\bar{\Lambda}_c^- \Delta_X(1600)^{++}$	( 4.6 $\pm$ 0.9 )	$\times 10^{-5}$ —
					$\bar{\Lambda}_c^- \Delta_X(2420)^{++}$	( 3.7 $\pm$ 0.8 )	$\times 10^{-5}$ —
					$(\bar{\Lambda}_c^- p)_s \pi^+$	[ <i>bbaa</i> ] ( 3.1 $\pm$ 0.7 )	$\times 10^{-5}$ —

# Meson Summary Table

$\Sigma_c(2520)^0 p$	$< 3$	$\times 10^{-6}$	CL=90%	1904
$\Sigma_c(2800)^0 p$	$(2.6 \pm 0.9)$	$\times 10^{-5}$		—
$\Lambda_c^- p \pi^+ \pi^0$	$(1.8 \pm 0.6)$	$\times 10^{-3}$		1935
$\Lambda_c^- p \pi^+ \pi^+ \pi^-$	$(2.2 \pm 0.7)$	$\times 10^{-3}$		1880
$\Lambda_c^- p \pi^+ \pi^+ \pi^- \pi^0$	$< 1.34$	%	CL=90%	1823
$\Lambda_c^+ \Lambda_c^- K^+$	$(6.9 \pm 2.2)$	$\times 10^{-4}$		—
$\Sigma_c(2455)^0 p$	$(2.9 \pm 0.7)$	$\times 10^{-5}$		1938
$\Sigma_c(2455)^0 p \pi^0$	$(3.5 \pm 1.1)$	$\times 10^{-4}$		1896
$\Sigma_c(2455)^0 p \pi^- \pi^+$	$(3.5 \pm 1.0)$	$\times 10^{-4}$		1845
$\Sigma_c(2455)^- p \pi^+ \pi^+$	$(2.34 \pm 0.20)$	$\times 10^{-4}$		1845
$\Lambda_c(2593)^- / \Lambda_c(2625)^- p \pi^+$	$< 1.9$	$\times 10^{-4}$	CL=90%	—
$\Xi_c^0 \Lambda_c^+ \rightarrow \Xi_c^- \Xi^+ \pi^-$	$(2.4 \pm 0.9)$	$\times 10^{-5}$	S=1.4	1144
$\Xi_c^0 \Lambda_c^+ \rightarrow \Xi_c^- \Lambda K^+ \pi^-$	$(2.1 \pm 0.9)$	$\times 10^{-5}$	S=1.5	1144

**Lepton Family number (LF) or Lepton number (L) or Baryon number (B) violating modes, or/and  $\Delta B = 1$  weak neutral current (BI) modes**

$\pi^+ \ell^+ \ell^-$	B1	$< 4.9$	$\times 10^{-8}$	CL=90%	2638
$\pi^+ e^+ e^-$	B1	$< 8.0$	$\times 10^{-8}$	CL=90%	2638
$\pi^+ \mu^+ \mu^-$	B1	$(1.79 \pm 0.23)$	$\times 10^{-8}$		2634
$\pi^+ \nu \bar{\nu}$	B1	$< 9.8$	$\times 10^{-5}$	CL=90%	2638
$K^+ \ell^+ \ell^-$	B1 [ttt]	$(4.51 \pm 0.23)$	$\times 10^{-7}$	S=1.1	2617
$K^+ e^+ e^-$	B1	$(5.5 \pm 0.7)$	$\times 10^{-7}$		2617
$K^+ \mu^+ \mu^-$	B1	$(4.43 \pm 0.24)$	$\times 10^{-7}$	S=1.2	2612
$K^+ \bar{\nu} \nu$	B1	$< 1.6$	$\times 10^{-5}$	CL=90%	2617
$\rho^+ \nu \bar{\nu}$	B1	$< 2.13$	$\times 10^{-4}$	CL=90%	2583
$K^*(892)^+ \ell^+ \ell^-$	B1 [ttt]	$(1.01 \pm 0.11)$	$\times 10^{-6}$	S=1.1	2564
$K^*(892)^+ e^+ e^-$	B1	$(1.55 \pm 0.40)$	$\times 10^{-6}$		2564
$K^*(892)^+ \mu^+ \mu^-$	B1	$(9.6 \pm 1.0)$	$\times 10^{-7}$		2560
$K^*(892)^+ \nu \bar{\nu}$	B1	$< 4.0$	$\times 10^{-5}$	CL=90%	2564
$K^+ \pi^+ \pi^- \mu^+ \mu^-$	B1	$(4.4 \pm 0.4)$	$\times 10^{-7}$		2593
$\phi K^+ \mu^+ \mu^-$	B1	$(7.9 \pm 2.1)$	$\times 10^{-8}$		2490
$\pi^+ e^+ \mu^-$	LF	$< 6.4$	$\times 10^{-3}$	CL=90%	2637
$\pi^+ e^- \mu^+$	LF	$< 6.4$	$\times 10^{-3}$	CL=90%	2637
$\pi^+ e^\pm \mu^\mp$	LF	$< 1.7$	$\times 10^{-7}$	CL=90%	2637
$\pi^+ e^+ \tau^-$	LF	$< 7.4$	$\times 10^{-5}$	CL=90%	2338
$\pi^+ e^- \tau^+$	LF	$< 2.0$	$\times 10^{-5}$	CL=90%	2338
$\pi^+ e^\pm \tau^\mp$	LF	$< 7.5$	$\times 10^{-5}$	CL=90%	2338
$\pi^+ \mu^+ \tau^-$	LF	$< 6.2$	$\times 10^{-5}$	CL=90%	2333
$\pi^+ \mu^- \tau^+$	LF	$< 4.5$	$\times 10^{-5}$	CL=90%	2333
$\pi^+ \mu^\pm \tau^\mp$	LF	$< 7.2$	$\times 10^{-5}$	CL=90%	2333
$K^+ e^+ \mu^-$	LF	$< 9.1$	$\times 10^{-8}$	CL=90%	2615
$K^+ e^- \mu^+$	LF	$< 1.3$	$\times 10^{-7}$	CL=90%	2615
$K^+ e^\pm \mu^\mp$	LF	$< 9.1$	$\times 10^{-8}$	CL=90%	2615
$K^+ e^+ \tau^-$	LF	$< 4.3$	$\times 10^{-5}$	CL=90%	2312
$K^+ e^- \tau^+$	LF	$< 1.5$	$\times 10^{-5}$	CL=90%	2312
$K^+ e^\pm \tau^\mp$	LF	$< 3.0$	$\times 10^{-5}$	CL=90%	2312
$K^+ \mu^+ \tau^-$	LF	$< 4.5$	$\times 10^{-5}$	CL=90%	2298
$K^+ \mu^- \tau^+$	LF	$< 2.8$	$\times 10^{-5}$	CL=90%	2298
$K^+ \mu^\pm \tau^\mp$	LF	$< 4.8$	$\times 10^{-5}$	CL=90%	2298
$K^*(892)^+ e^+ \mu^-$	LF	$< 1.3$	$\times 10^{-6}$	CL=90%	2563
$K^*(892)^+ e^- \mu^+$	LF	$< 9.9$	$\times 10^{-7}$	CL=90%	2563
$K^*(892)^+ e^\pm \mu^\mp$	LF	$< 1.4$	$\times 10^{-6}$	CL=90%	2563
$\pi^- e^+ e^+$	L	$< 2.3$	$\times 10^{-8}$	CL=90%	2638
$\pi^- \mu^+ \mu^+$	L	$< 4.0$	$\times 10^{-9}$	CL=95%	2634
$\pi^- e^+ \mu^+$	L	$< 1.5$	$\times 10^{-7}$	CL=90%	2637
$\rho^- e^+ e^+$	L	$< 1.7$	$\times 10^{-7}$	CL=90%	2583
$\rho^- \mu^+ \mu^+$	L	$< 4.2$	$\times 10^{-7}$	CL=90%	2578
$\rho^- e^+ \mu^+$	L	$< 4.7$	$\times 10^{-7}$	CL=90%	2582
$K^- e^+ e^+$	L	$< 3.0$	$\times 10^{-8}$	CL=90%	2617
$K^- \mu^+ \mu^+$	L	$< 4.1$	$\times 10^{-8}$	CL=90%	2612
$K^- e^+ \mu^+$	L	$< 1.6$	$\times 10^{-7}$	CL=90%	2615
$K^*(892)^- e^+ e^+$	L	$< 4.0$	$\times 10^{-7}$	CL=90%	2564
$K^*(892)^- \mu^+ \mu^+$	L	$< 5.9$	$\times 10^{-7}$	CL=90%	2560
$K^*(892)^- e^+ \mu^+$	L	$< 3.0$	$\times 10^{-7}$	CL=90%	2563
$D^- e^+ e^+$	L	$< 2.6$	$\times 10^{-6}$	CL=90%	2309
$D^- e^+ \mu^+$	L	$< 1.8$	$\times 10^{-6}$	CL=90%	2307
$D^- \mu^+ \mu^+$	L	$< 6.9$	$\times 10^{-7}$	CL=95%	2303
$D^* \mu^+ \mu^+$	L	$< 2.4$	$\times 10^{-6}$	CL=95%	2251
$D_s^- \mu^+ \mu^+$	L	$< 5.8$	$\times 10^{-7}$	CL=95%	2267
$\bar{D}^0 \pi^- \mu^+ \mu^+$	L	$< 1.5$	$\times 10^{-6}$	CL=95%	2295
$\Lambda^0 \mu^+$	L,B	$< 6$	$\times 10^{-8}$	CL=90%	—
$\Lambda^0 e^+$	L,B	$< 3.2$	$\times 10^{-8}$	CL=90%	—
$\bar{\Lambda}^0 \mu^+$	L,B	$< 6$	$\times 10^{-8}$	CL=90%	—
$\bar{\Lambda}^0 e^+$	L,B	$< 8$	$\times 10^{-8}$	CL=90%	—

**$B^0$**

$$I(J^P) = \frac{1}{2}(0^-)$$

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

$$\text{Mass } m_{B^0} = 5279.62 \pm 0.15 \text{ MeV} \quad (S = 1.1)$$

$$m_{B^0} - m_{B^\pm} = 0.31 \pm 0.06 \text{ MeV}$$

$$\text{Mean life } \tau_{B^0} = (1.520 \pm 0.004) \times 10^{-12} \text{ s}$$

$$c\tau = 455.7 \text{ } \mu\text{m}$$

$$\tau_{B^+}/\tau_{B^0} = 1.076 \pm 0.004 \quad (\text{direct measurements})$$

## $B^0$ - $\bar{B}^0$ mixing parameters

$$\chi_d = 0.1875 \pm 0.0017$$

$$\Delta m_{B^0} = m_{B_H^0} - m_{B_L^0} = (0.5096 \pm 0.0034) \times 10^{12} \text{ } \hbar \text{ s}^{-1} \\ = (3.354 \pm 0.022) \times 10^{-10} \text{ MeV}$$

$$x_d = \Delta m_{B^0}/\Gamma_{B^0} = 0.775 \pm 0.006$$

$$\text{Re}(\lambda_{CP} / |\lambda_{CP}|) \text{ Re}(z) = 0.01 \pm 0.05$$

$$\Delta\Gamma \text{ Re}(z) = -0.007 \pm 0.004$$

$$\text{Re}(z) = (2 \pm 5) \times 10^{-2}$$

$$\text{Im}(z) = (-0.8 \pm 0.4) \times 10^{-2}$$

## CP violation parameters

$$\text{Re}(\epsilon_{B^0})/(1+|\epsilon_{B^0}|^2) = (-0.4 \pm 0.4) \times 10^{-3}$$

$$A_{T/CP} = 0.005 \pm 0.018$$

$$A_{CP}(B^0 \rightarrow D^*(2010)^+ D^-) = 0.037 \pm 0.034$$

$$A_{CP}(B^0 \rightarrow [K^+ K^-]_D K^*(892)^0) = -0.20 \pm 0.15$$

$$A_{CP}(B^0 \rightarrow [K^+ \pi^-]_D K^*(892)^0) = -0.03 \pm 0.04$$

$$R_d^+ = \Gamma(B^0 \rightarrow [\pi^+ K^-]_D K^{*0}) / \Gamma(B^0 \rightarrow [\pi^- K^+]_D K^{*0}) = \\ 0.06 \pm 0.032$$

$$R_d^- = \Gamma(\bar{B}^0 \rightarrow [\pi^- K^+]_D K^{*0}) / \Gamma(\bar{B}^0 \rightarrow [\pi^+ K^-]_D K^{*0}) = \\ 0.06 \pm 0.032$$

$$A_{CP}(B^0 \rightarrow [\pi^+ \pi^-]_D K^*(892)^0) = -0.09 \pm 0.22$$

$$A_{CP}(B^0 \rightarrow K^+ \pi^-) = -0.082 \pm 0.006$$

$$A_{CP}(B^0 \rightarrow \eta' K^*(892)^0) = -0.07 \pm 0.18$$

$$A_{CP}(B^0 \rightarrow \eta' K_0^*(1430)^0) = -0.19 \pm 0.17$$

$$A_{CP}(B^0 \rightarrow \eta' K_2^*(1430)^0) = 0.14 \pm 0.18$$

$$A_{CP}(B^0 \rightarrow \eta K^*(892)^0) = 0.19 \pm 0.05$$

$$A_{CP}(B^0 \rightarrow \eta K_0^*(1430)^0) = 0.06 \pm 0.13$$

$$A_{CP}(B^0 \rightarrow \eta K_2^*(1430)^0) = -0.07 \pm 0.19$$

$$A_{CP}(B^0 \rightarrow b_1 K^+) = -0.07 \pm 0.12$$

$$A_{CP}(B^0 \rightarrow \omega K^{*0}) = 0.45 \pm 0.25$$

$$A_{CP}(B^0 \rightarrow \omega (K\pi)_0^{*0}) = -0.07 \pm 0.09$$

$$A_{CP}(B^0 \rightarrow \omega K_2^*(1430)^0) = -0.37 \pm 0.17$$

$$A_{CP}(B^0 \rightarrow K^+ \pi^- \pi^0) = (0 \pm 6) \times 10^{-2}$$

$$A_{CP}(B^0 \rightarrow \rho^- K^+) = 0.20 \pm 0.11$$

$$A_{CP}(B^0 \rightarrow \rho(1450)^- K^+) = -0.10 \pm 0.33$$

$$A_{CP}(B^0 \rightarrow \rho(1700)^- K^+) = -0.4 \pm 0.6$$

$$A_{CP}(B^0 \rightarrow K^+ \pi^- \pi^0 \text{ nonresonant}) = 0.10 \pm 0.18$$

$$A_{CP}(B^0 \rightarrow K^0 \pi^+ \pi^-) = -0.01 \pm 0.05$$

$$A_{CP}(B^0 \rightarrow K^*(892)^+ \pi^-) = -0.22 \pm 0.06$$

$$A_{CP}(B^0 \rightarrow (K\pi)_0^{*+} \pi^-) = 0.09 \pm 0.07$$

$$A_{CP}(B^0 \rightarrow (K\pi)_0^{*0} \pi^0) = -0.15 \pm 0.11$$

$$A_{CP}(B^0 \rightarrow K^{*0} \pi^0) = -0.15 \pm 0.13$$

$$A_{CP}(B^0 \rightarrow K^*(892)^0 \pi^+ \pi^-) = 0.07 \pm 0.05$$

$$A_{CP}(B^0 \rightarrow K^*(892)^0 \rho^0) = -0.06 \pm 0.09$$

$$A_{CP}(B^0 \rightarrow K^{*0} f_0(980)) = 0.07 \pm 0.10$$

$$A_{CP}(B^0 \rightarrow K^{*+} \rho^-) = 0.21 \pm 0.15$$

$$A_{CP}(B^0 \rightarrow K^*(892)^0 K^+ K^-) = 0.01 \pm 0.05$$

$$A_{CP}(B^0 \rightarrow a_1^- K^+) = -0.16 \pm 0.12$$

$$A_{CP}(B^0 \rightarrow K^0 K^0) = -0.6 \pm 0.7$$

$$A_{CP}(B^0 \rightarrow K^*(892)^0 \phi) = 0.00 \pm 0.04$$

$$A_{CP}(B^0 \rightarrow K^*(892)^0 K^- \pi^+) = 0.2 \pm 0.4$$

$$A_{CP}(B^0 \rightarrow \phi (K\pi)_0^{*0}) = 0.12 \pm 0.08$$

$$A_{CP}(B^0 \rightarrow \phi K_2^*(1430)^0) = -0.11 \pm 0.10$$

$$A_{CP}(B^0 \rightarrow K^*(892)^0 \gamma) = -0.002 \pm 0.015$$

$$A_{CP}(B^0 \rightarrow K_2^*(1430)^0 \gamma) = -0.08 \pm 0.15$$

$$A_{CP}(B^0 \rightarrow \rho^+ \pi^-) = 0.13 \pm 0.06 \quad (S = 1.1)$$

$$A_{CP}(B^0 \rightarrow \rho^- \pi^+) = -0.08 \pm 0.08$$

$$A_{CP}(B^0 \rightarrow a_1(1260)^\pm \pi^\mp) = -0.07 \pm 0.06$$

$$A_{CP}(B^0 \rightarrow b_1^- \pi^+) = -0.05 \pm 0.10$$

$$A_{CP}(B^0 \rightarrow p \bar{p} K^*(892)^0) = 0.05 \pm 0.12$$

$$A_{CP}(B^0 \rightarrow p \bar{\Lambda} \pi^-) = 0.04 \pm 0.07$$

$$A_{CP}(B^0 \rightarrow K^{*0} \ell^+ \ell^-) = -0.05 \pm 0.10$$

$$A_{CP}(B^0 \rightarrow K^{*0} e^+ e^-) = -0.21 \pm 0.19$$

## Meson Summary Table

$A_{CP}(B^0 \rightarrow K^{*0} \mu^+ \mu^-) = -0.034 \pm 0.024$	$S(B^0 \rightarrow K_S^0 \rho^0 \gamma) = 0.11 \pm 0.34$
$C_{D^{*-} D^+}(B^0 \rightarrow D^*(2010)^- D^+) = -0.01 \pm 0.11$	$C(B^0 \rightarrow \rho^0 \gamma) = 0.4 \pm 0.5$
$S_{D^{*-} D^+}(B^0 \rightarrow D^*(2010)^- D^+) = -0.72 \pm 0.15$	$S(B^0 \rightarrow \rho^0 \gamma) = -0.8 \pm 0.7$
$C_{D^{*+} D^-}(B^0 \rightarrow D^*(2010)^+ D^-) = 0.00 \pm 0.13 \quad (S = 1.3)$	$C_{\pi\pi}(B^0 \rightarrow \pi^+ \pi^-) = -0.31 \pm 0.05$
$S_{D^{*+} D^-}(B^0 \rightarrow D^*(2010)^+ D^-) = -0.73 \pm 0.14$	$S_{\pi\pi}(B^0 \rightarrow \pi^+ \pi^-) = -0.67 \pm 0.06$
$C_{D^{*+} D^{*-}}(B^0 \rightarrow D^{*+} D^{*-}) = 0.01 \pm 0.09 \quad (S = 1.6)$	$C_{\pi^0 \pi^0}(B^0 \rightarrow \pi^0 \pi^0) = -0.43 \pm 0.24$
$S_{D^{*+} D^{*-}}(B^0 \rightarrow D^{*+} D^{*-}) = -0.59 \pm 0.14 \quad (S = 1.8)$	$C_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-) = -0.03 \pm 0.07 \quad (S = 1.2)$
$C_+(B^0 \rightarrow D^{*+} D^{*-}) = 0.00 \pm 0.10 \quad (S = 1.6)$	$S_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-) = 0.05 \pm 0.07$
$S_+(B^0 \rightarrow D^{*+} D^{*-}) = -0.73 \pm 0.09$	$\Delta C_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-) = 0.27 \pm 0.06$
$C_-(B^0 \rightarrow D^{*+} D^{*-}) = 0.19 \pm 0.31$	$\Delta S_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-) = 0.01 \pm 0.08$
$S_-(B^0 \rightarrow D^{*+} D^{*-}) = 0.1 \pm 1.6 \quad (S = 3.5)$	$C_{\rho^0 \pi^0}(B^0 \rightarrow \rho^0 \pi^0) = 0.27 \pm 0.24$
$C(B^0 \rightarrow D^*(2010)^+ D^*(2010)^- K_S^0) = 0.01 \pm 0.29$	$S_{\rho^0 \pi^0}(B^0 \rightarrow \rho^0 \pi^0) = -0.23 \pm 0.34$
$S(B^0 \rightarrow D^*(2010)^+ D^*(2010)^- K_S^0) = 0.1 \pm 0.4$	$C_{a_1 \pi}(B^0 \rightarrow a_1(1260)^+ \pi^-) = -0.05 \pm 0.11$
$C_{D^+ D^-}(B^0 \rightarrow D^+ D^-) = -0.46 \pm 0.21 \quad (S = 1.8)$	$S_{a_1 \pi}(B^0 \rightarrow a_1(1260)^+ \pi^-) = -0.2 \pm 0.4 \quad (S = 3.2)$
$S_{D^+ D^-}(B^0 \rightarrow D^+ D^-) = -0.99 \pm 0.17$	$\Delta C_{a_1 \pi}(B^0 \rightarrow a_1(1260)^+ \pi^-) = 0.43 \pm 0.14 \quad (S = 1.3)$
$C_{J/\psi(1S) \pi^0}(B^0 \rightarrow J/\psi(1S) \pi^0) = -0.13 \pm 0.13$	$\Delta S_{a_1 \pi}(B^0 \rightarrow a_1(1260)^+ \pi^-) = -0.11 \pm 0.12$
$S_{J/\psi(1S) \pi^0}(B^0 \rightarrow J/\psi(1S) \pi^0) = -0.94 \pm 0.29 \quad (S = 1.9)$	$C(B^0 \rightarrow b_1^- K^+) = -0.22 \pm 0.24$
$C(B^0 \rightarrow J/\psi(1S) \rho^0) = -0.06 \pm 0.06$	$\Delta C(B^0 \rightarrow b_1^- K^+) = -1.04 \pm 0.24$
$S(B^0 \rightarrow J/\psi(1S) \rho^0) = -0.66 \pm 0.16$	$C_{\rho^0 \rho^0}(B^0 \rightarrow \rho^0 \rho^0) = 0.2 \pm 0.9$
$C_{D_{CP}^{(*)} h^0}(B^0 \rightarrow D_{CP}^{(*)} h^0) = -0.02 \pm 0.08$	$S_{\rho^0 \rho^0}(B^0 \rightarrow \rho^0 \rho^0) = 0.3 \pm 0.7$
$S_{D_{CP}^{(*)} h^0}(B^0 \rightarrow D_{CP}^{(*)} h^0) = -0.66 \pm 0.12$	$C_{\rho\rho}(B^0 \rightarrow \rho^+ \rho^-) = 0.00 \pm 0.09$
$C_{K^0 \pi^0}(B^0 \rightarrow K^0 \pi^0) = 0.00 \pm 0.13 \quad (S = 1.4)$	$S_{\rho\rho}(B^0 \rightarrow \rho^+ \rho^-) = -0.14 \pm 0.13$
$S_{K^0 \pi^0}(B^0 \rightarrow K^0 \pi^0) = 0.58 \pm 0.17$	$ \lambda (B^0 \rightarrow J/\psi K^*(892)^0) < 0.25, \text{CL} = 95\%$
$C_{\eta'(958) K_S^0}(B^0 \rightarrow \eta'(958) K_S^0) = -0.04 \pm 0.20 \quad (S = 2.5)$	$\cos 2\beta(B^0 \rightarrow J/\psi K^*(892)^0) = 1.7 \pm 0.7 \quad (S = 1.6)$
$S_{\eta'(958) K_S^0}(B^0 \rightarrow \eta'(958) K_S^0) = 0.43 \pm 0.17 \quad (S = 1.5)$	$\cos 2\beta(B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0) = 1.0 \pm 0.6 \quad (S = 1.8)$
$C_{\eta' K^0}(B^0 \rightarrow \eta' K^0) = -0.06 \pm 0.04$	$(S_+ + S_-)/2(B^0 \rightarrow D^{*-} \pi^+) = -0.039 \pm 0.011$
$S_{\eta' K^0}(B^0 \rightarrow \eta' K^0) = 0.63 \pm 0.06$	$(S_- - S_+)/2(B^0 \rightarrow D^{*-} \pi^+) = -0.009 \pm 0.015$
$C_{\omega K_S^0}(B^0 \rightarrow \omega K_S^0) = 0.0 \pm 0.4 \quad (S = 3.0)$	$(S_+ + S_-)/2(B^0 \rightarrow D^- \pi^+) = -0.046 \pm 0.023$
$S_{\omega K_S^0}(B^0 \rightarrow \omega K_S^0) = 0.70 \pm 0.21$	$(S_- - S_+)/2(B^0 \rightarrow D^- \pi^+) = -0.022 \pm 0.021$
$C(B^0 \rightarrow K_S^0 \pi^0 \pi^0) = 0.2 \pm 0.5$	$(S_+ + S_-)/2(B^0 \rightarrow D^- \rho^+) = -0.024 \pm 0.032$
$S(B^0 \rightarrow K_S^0 \pi^0 \pi^0) = 0.7 \pm 0.7$	$(S_- - S_+)/2(B^0 \rightarrow D^- \rho^+) = -0.10 \pm 0.06$
$C_{\rho^0 K_S^0}(B^0 \rightarrow \rho^0 K_S^0) = -0.04 \pm 0.20$	$C_{\eta_c K_S^0}(B^0 \rightarrow \eta_c K_S^0) = 0.08 \pm 0.13$
$S_{\rho^0 K_S^0}(B^0 \rightarrow \rho^0 K_S^0) = 0.50 \pm 0.17$	$S_{\eta_c K_S^0}(B^0 \rightarrow \eta_c K_S^0) = 0.93 \pm 0.17$
$C_{f_0 K_S^0}(B^0 \rightarrow f_0(980) K_S^0) = 0.29 \pm 0.20$	$C_{c\bar{c}K^{(*)0}}(B^0 \rightarrow c\bar{c}K^{(*)0}) = (0.5 \pm 1.7) \times 10^{-2}$
$S_{f_0 K_S^0}(B^0 \rightarrow f_0(980) K_S^0) = -0.50 \pm 0.16$	$\sin(2\beta) = 0.679 \pm 0.020$
$S_{f_2 K_S^0}(B^0 \rightarrow f_2(1270) K_S^0) = -0.5 \pm 0.5$	$C_{J/\psi(nS) K^0}(B^0 \rightarrow J/\psi(nS) K^0) = (0.5 \pm 2.0) \times 10^{-2}$
$C_{f_2 K_S^0}(B^0 \rightarrow f_2(1270) K_S^0) = 0.3 \pm 0.4$	$S_{J/\psi(nS) K^0}(B^0 \rightarrow J/\psi(nS) K^0) = 0.676 \pm 0.021$
$S_{f_x K_S^0}(B^0 \rightarrow f_x(1300) K_S^0) = -0.2 \pm 0.5$	$C_{J/\psi K^{*0}}(B^0 \rightarrow J/\psi K^{*0}) = 0.03 \pm 0.10$
$C_{f_x K_S^0}(B^0 \rightarrow f_x(1300) K_S^0) = 0.13 \pm 0.35$	$S_{J/\psi K^{*0}}(B^0 \rightarrow J/\psi K^{*0}) = 0.60 \pm 0.25$
$S_{K^0 \pi^+ \pi^-}(B^0 \rightarrow K^0 \pi^+ \pi^- \text{ nonresonant}) = -0.01 \pm 0.33$	$C_{\chi_{c0} K_S^0}(B^0 \rightarrow \chi_{c0} K_S^0) = -0.3 \pm 0.5$
$C_{K^0 \pi^+ \pi^-}(B^0 \rightarrow K^0 \pi^+ \pi^- \text{ nonresonant}) = 0.01 \pm 0.26$	$S_{\chi_{c0} K_S^0}(B^0 \rightarrow \chi_{c0} K_S^0) = -0.7 \pm 0.5$
$C_{K_S^0 K_S^0}(B^0 \rightarrow K_S^0 K_S^0) = 0.0 \pm 0.4 \quad (S = 1.4)$	$C_{\chi_{c1} K_S^0}(B^0 \rightarrow \chi_{c1} K_S^0) = 0.06 \pm 0.07$
$S_{K_S^0 K_S^0}(B^0 \rightarrow K_S^0 K_S^0) = -0.8 \pm 0.5$	$S_{\chi_{c1} K_S^0}(B^0 \rightarrow \chi_{c1} K_S^0) = 0.63 \pm 0.10$
$C_{K^+ K^- K_S^0}(B^0 \rightarrow K^+ K^- K_S^0 \text{ nonresonant}) = 0.06 \pm 0.08$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow \phi K^0) = 0.22 \pm 0.30$
$S_{K^+ K^- K_S^0}(B^0 \rightarrow K^+ K^- K_S^0 \text{ nonresonant}) = -0.66 \pm 0.11$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow \phi K_0^0(1430)^0) = 0.97 \pm 0.03$
$C_{K^+ K^- K_S^0}(B^0 \rightarrow K^+ K^- K_S^0 \text{ inclusive}) = 0.01 \pm 0.09$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow K^+ K^- K_S^0) = 0.77 \pm 0.13$
$S_{K^+ K^- K_S^0}(B^0 \rightarrow K^+ K^- K_S^0 \text{ inclusive}) = -0.65 \pm 0.12$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0) = 0.45 \pm 0.28$
$C_{\phi K_S^0}(B^0 \rightarrow \phi K_S^0) = 0.01 \pm 0.14$	$2\beta_{\text{eff}}(B^0 \rightarrow J/\psi \rho^0) = (42 \pm 11)^\circ$
$S_{\phi K_S^0}(B^0 \rightarrow \phi K_S^0) = 0.59 \pm 0.14$	$ \lambda (B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0) = 1.01 \pm 0.08$
$C_{K_S K_S K_S}(B^0 \rightarrow K_S K_S K_S) = -0.23 \pm 0.14$	$ \sin(2\beta + \gamma)  > 0.40, \text{CL} = 90\%$
$S_{K_S K_S K_S}(B^0 \rightarrow K_S K_S K_S) = -0.5 \pm 0.6 \quad (S = 3.0)$	$2\beta + \gamma = (83 \pm 60)^\circ$
$C_{K_S^0 \pi^0 \gamma}(B^0 \rightarrow K_S^0 \pi^0 \gamma) = 0.36 \pm 0.33$	$\gamma(B^0 \rightarrow D^0 K^{*0}) = (162 \pm 60)^\circ$
$S_{K_S^0 \pi^0 \gamma}(B^0 \rightarrow K_S^0 \pi^0 \gamma) = -0.8 \pm 0.6$	$\alpha = (93 \pm 5)^\circ$
$C_{K^{*0} \gamma}(B^0 \rightarrow K^*(892)^0 \gamma) = -0.04 \pm 0.16 \quad (S = 1.2)$	
$S_{K^{*0} \gamma}(B^0 \rightarrow K^*(892)^0 \gamma) = -0.15 \pm 0.22$	
$C_{\eta K^0 \gamma}(B^0 \rightarrow \eta K^0 \gamma) = -0.3 \pm 0.4$	
$S_{\eta K^0 \gamma}(B^0 \rightarrow \eta K^0 \gamma) = -0.2 \pm 0.5$	
$C_{K^0 \phi \gamma}(B^0 \rightarrow K^0 \phi \gamma) = -0.3 \pm 0.6$	
$S_{K^0 \phi \gamma}(B^0 \rightarrow K^0 \phi \gamma) = 0.7 \pm 0.7$	
$C(B^0 \rightarrow K_S^0 \rho^0 \gamma) = -0.05 \pm 0.19$	

# Meson Summary Table

$\overline{B}^0$  modes are charge conjugates of the modes below. Reactions indicate the weak decay vertex and do not include mixing. Modes which do not identify the charge state of the  $B$  are listed in the  $B^\pm/B^0$  ADMIXTURE section.

The branching fractions listed below assume 50%  $B^0 \overline{B}^0$  and 50%  $B^+ B^-$  production at the  $\Upsilon(4S)$ . We have attempted to bring older measurements up to date by rescaling their assumed  $\Upsilon(4S)$  production ratio to 50:50 and their assumed  $D$ ,  $D_s$ ,  $D^*$ , and  $\psi$  branching ratios to current values whenever this would affect our averages and best limits significantly.

Indentation is used to indicate a subchannel of a previous reaction. All resonant subchannels have been corrected for resonance branching fractions to the final state so the sum of the subchannel branching fractions can exceed that of the final state.

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm$  anything, the values usually are multiplicities, not branching fractions. They can be greater than one.

$B^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$\ell^+ \nu_\ell$ anything	[ <i>ttt</i> ] ( 10.33 $\pm$ 0.28 ) %	—	—
$e^+ \nu_e X_c$	( 10.1 $\pm$ 0.4 ) %	—	—
$D \ell^+ \nu_\ell$ anything	( 9.2 $\pm$ 0.8 ) %	—	—
$D^- \ell^+ \nu_\ell$	[ <i>ttt</i> ] ( 2.19 $\pm$ 0.12 ) %	2309	—
$D^- \tau^+ \nu_\tau$	( 1.03 $\pm$ 0.22 ) %	1909	—
$D^*(2010)^- \ell^+ \nu_\ell$	[ <i>ttt</i> ] ( 4.93 $\pm$ 0.11 ) %	2257	—
$D^*(2010)^- \tau^+ \nu_\tau$	( 1.78 $\pm$ 0.17 ) %	S=1.1 1837	—
$\overline{D}^0 \pi^- \ell^+ \nu_\ell$	( 4.3 $\pm$ 0.6 ) $\times 10^{-3}$	2308	—
$D_0^*(2400)^- \ell^+ \nu_\ell, D_0^{*-} \rightarrow \overline{D}^0 \pi^-$	( 3.0 $\pm$ 1.2 ) $\times 10^{-3}$	S=1.8 —	—
$D_2^*(2460)^- \ell^+ \nu_\ell, D_2^{*-} \rightarrow \overline{D}^0 \pi^-$	( 1.21 $\pm$ 0.33 ) $\times 10^{-3}$	S=1.8 2065	—
$\overline{D}^{(*)} n \pi \ell^+ \nu_\ell (n \geq 1)$	( 2.3 $\pm$ 0.5 ) %	—	—
$\overline{D}^{*0} \pi^- \ell^+ \nu_\ell$	( 4.9 $\pm$ 0.8 ) $\times 10^{-3}$	2256	—
$D_1(2420)^- \ell^+ \nu_\ell, D_1^- \rightarrow \overline{D}^{*0} \pi^-$	( 2.80 $\pm$ 0.28 ) $\times 10^{-3}$	—	—
$D_1'(2430)^- \ell^+ \nu_\ell, D_1'^- \rightarrow \overline{D}^{*0} \pi^-$	( 3.1 $\pm$ 0.9 ) $\times 10^{-3}$	—	—
$D_2^*(2460)^- \ell^+ \nu_\ell, D_2^{*-} \rightarrow \overline{D}^{*0} \pi^-$	( 6.8 $\pm$ 1.2 ) $\times 10^{-4}$	2065	—
$D^- \pi^+ \pi^- \ell^+ \nu_\ell$	( 1.3 $\pm$ 0.5 ) $\times 10^{-3}$	2299	—
$D^{*-} \pi^+ \pi^- \ell^+ \nu_\ell$	( 1.4 $\pm$ 0.5 ) $\times 10^{-3}$	2247	—
$\rho^- \ell^+ \nu_\ell$	[ <i>ttt</i> ] ( 2.94 $\pm$ 0.21 ) $\times 10^{-4}$	2583	—
$\pi^- \ell^+ \nu_\ell$	[ <i>ttt</i> ] ( 1.45 $\pm$ 0.05 ) $\times 10^{-4}$	2638	—
$\pi^- \tau^+ \nu_\tau$	< 2.5 $\times 10^{-4}$	CL=90%	2338
<b>Inclusive modes</b>			
$K^\pm$ anything	( 78 $\pm$ 8 ) %	—	—
$D^0 X$	( 8.1 $\pm$ 1.5 ) %	—	—
$\overline{D}^0 X$	( 47.4 $\pm$ 2.8 ) %	—	—
$D^+ X$	< 3.9 %	CL=90%	—
$D^- X$	( 36.9 $\pm$ 3.3 ) %	—	—
$D_s^+ X$	( 10.3 $\pm$ 2.1 ) %	—	—
$D_s^- X$	< 2.6 %	CL=90%	—
$\Lambda_c^+ X$	< 3.1 %	CL=90%	—
$\overline{\Lambda}_c^- X$	( 5.0 $\pm$ 1.5 ) %	—	—
$\overline{c} X$	( 95 $\pm$ 5 ) %	—	—
$c X$	( 24.6 $\pm$ 3.1 ) %	—	—
$\overline{c} c X$	( 119 $\pm$ 6 ) %	—	—
<b><math>D</math>, <math>D^*</math>, or <math>D_s</math> modes</b>			
$D^- \pi^+$	( 2.52 $\pm$ 0.13 ) $\times 10^{-3}$	S=1.1 2306	—
$D^- \rho^+$	( 7.5 $\pm$ 1.2 ) $\times 10^{-3}$	2235	—
$D^- K^0 \pi^+$	( 4.9 $\pm$ 0.9 ) $\times 10^{-4}$	2259	—
$D^- K^*(892)^+$	( 4.5 $\pm$ 0.7 ) $\times 10^{-4}$	2211	—
$D^- \omega \pi^+$	( 2.8 $\pm$ 0.6 ) $\times 10^{-3}$	2204	—
$D^- K^+$	( 1.86 $\pm$ 0.20 ) $\times 10^{-4}$	2279	—
$D^- K^+ \pi^+ \pi^-$	( 3.5 $\pm$ 0.8 ) $\times 10^{-4}$	2236	—
$D^- K^+ \overline{K}^0$	< 3.1 $\times 10^{-4}$	CL=90%	2188
$D^- K^+ \overline{K}^*(892)^0$	( 8.8 $\pm$ 1.9 ) $\times 10^{-4}$	2070	—
$\overline{D}^0 \pi^+ \pi^-$	( 8.8 $\pm$ 0.5 ) $\times 10^{-4}$	2301	—
$D^*(2010)^- \pi^+$	( 2.74 $\pm$ 0.13 ) $\times 10^{-3}$	2255	—
$\overline{D}^0 K^+ K^-$	( 4.9 $\pm$ 1.2 ) $\times 10^{-5}$	2191	—
$D^- \pi^+ \pi^+ \pi^-$	( 6.0 $\pm$ 0.7 ) $\times 10^{-3}$	S=1.1 2287	—
( $D^- \pi^+ \pi^+ \pi^-$ ) nonresonant	( 3.9 $\pm$ 1.9 ) $\times 10^{-3}$	2287	—
$D^- \pi^+ \rho^0$	( 1.1 $\pm$ 1.0 ) $\times 10^{-3}$	2206	—
$D^- a_1(1260)^+$	( 6.0 $\pm$ 3.3 ) $\times 10^{-3}$	2121	—
$D^*(2010)^- \pi^+ \pi^0$	( 1.5 $\pm$ 0.5 ) %	2248	—

$D^*(2010)^- \rho^+$	( 2.2 $\pm$ 1.8 ) $\times 10^{-3}$	S=5.2	2180
$D^*(2010)^- K^+$	( 2.12 $\pm$ 0.15 ) $\times 10^{-4}$	—	2226
$D^*(2010)^- K^0 \pi^+$	( 3.0 $\pm$ 0.8 ) $\times 10^{-4}$	—	2205
$D^*(2010)^- K^*(892)^+$	( 3.3 $\pm$ 0.6 ) $\times 10^{-4}$	—	2155
$D^*(2010)^- K^+ \overline{K}^0$	< 4.7 $\times 10^{-4}$	CL=90%	2131
$D^*(2010)^- K^+ \overline{K}^*(892)^0$	( 1.29 $\pm$ 0.33 ) $\times 10^{-3}$	—	2007
$D^*(2010)^- \pi^+ \pi^+ \pi^-$	( 7.0 $\pm$ 0.8 ) $\times 10^{-3}$	S=1.3	2235
( $D^*(2010)^- \pi^+ \pi^+ \pi^-$ ) non-resonant	( 0.0 $\pm$ 2.5 ) $\times 10^{-3}$	—	2235
$D^*(2010)^- \pi^+ \rho^0$	( 5.7 $\pm$ 3.2 ) $\times 10^{-3}$	—	2150
$D^*(2010)^- a_1(1260)^+$	( 1.30 $\pm$ 0.27 ) %	—	2061
$\overline{D}_1(2420)^0 \pi^- \pi^+, \overline{D}_1^0 \rightarrow D^{*-} \pi^+$	( 1.4 $\pm$ 0.4 ) $\times 10^{-4}$	—	—
$D^*(2010)^- K^+ \pi^- \pi^+$	( 4.5 $\pm$ 0.7 ) $\times 10^{-4}$	—	2181
$D^*(2010)^- \pi^+ \pi^+ \pi^- \pi^0$	( 1.76 $\pm$ 0.27 ) %	—	2218
$D^{*-} 3\pi^+ 2\pi^-$	( 4.7 $\pm$ 0.9 ) $\times 10^{-3}$	—	2195
$\overline{D}^*(2010)^- \omega \pi^+$	( 2.46 $\pm$ 0.18 ) $\times 10^{-3}$	S=1.2	2148
$D_1(2430)^0 \omega, D_1^0 \rightarrow D^{*-} \pi^+$	( 2.7 $\pm$ 0.8 ) $\times 10^{-4}$	—	1992
$\overline{D}^{*-} \rho(1450)^+$	( 1.07 $\pm$ 0.40 ) $\times 10^{-3}$	—	—
$\overline{D}_1(2420)^0 \omega$	( 7.0 $\pm$ 2.2 ) $\times 10^{-5}$	—	1995
$\overline{D}_2^*(2460)^0 \omega$	( 4.0 $\pm$ 1.4 ) $\times 10^{-5}$	—	1975
$\overline{D}^{*-} b_1(1235)^-, b_1^- \rightarrow \omega \pi^-$	< 7 $\times 10^{-5}$	CL=90%	—
$\overline{D}^{*-} \pi^+$	[ <i>yyy</i> ] ( 1.9 $\pm$ 0.9 ) $\times 10^{-3}$	—	—
$D_1(2420)^- \pi^+, D_1^- \rightarrow D^- \pi^+ \pi^-$	( 9.9 $\pm$ 2.5 ) $\times 10^{-5}$	—	—
$D_1(2420)^- \pi^+, D_1^- \rightarrow D^{*-} \pi^+ \pi^-$	< 3.3 $\times 10^{-5}$	CL=90%	—
$\overline{D}_2^*(2460)^- \pi^+, (D_2^*)^- \rightarrow D^0 \pi^-$	( 2.38 $\pm$ 0.16 ) $\times 10^{-4}$	—	2062
$\overline{D}_0^*(2400)^- \pi^+, (D_0^*)^- \rightarrow D^0 \pi^-$	( 7.6 $\pm$ 0.8 ) $\times 10^{-5}$	—	2090
$D_2^*(2460)^- \pi^+, (D_2^*)^- \rightarrow D^{*-} \pi^+ \pi^-$	< 2.4 $\times 10^{-5}$	CL=90%	—
$\overline{D}_2^*(2460)^- \rho^+$	< 4.9 $\times 10^{-3}$	CL=90%	1974
$D^0 \overline{D}^0$	( 1.4 $\pm$ 0.7 ) $\times 10^{-5}$	—	1868
$D^{*0} \overline{D}^0$	< 2.9 $\times 10^{-4}$	CL=90%	1794
$D^- D^+$	( 2.11 $\pm$ 0.18 ) $\times 10^{-4}$	—	1864
$D^\pm D^{*\mp}$ ( $CP$ -averaged)	( 6.1 $\pm$ 0.6 ) $\times 10^{-4}$	—	—
$D^- D_s^+$	( 7.2 $\pm$ 0.8 ) $\times 10^{-3}$	—	1813
$D^*(2010)^- D_s^+$	( 8.0 $\pm$ 1.1 ) $\times 10^{-3}$	—	1735
$D^- D_s^{*+}$	( 7.4 $\pm$ 1.6 ) $\times 10^{-3}$	—	1732
$D^*(2010)^- D_s^{*+}$	( 1.77 $\pm$ 0.14 ) %	—	1649
$D_{s0}(2317)^- K^+, D_{s0}^- \rightarrow D_s^- \pi^0$	( 4.2 $\pm$ 1.4 ) $\times 10^{-5}$	—	2097
$D_{s0}(2317)^- \pi^+, D_{s0}^- \rightarrow D_s^- \pi^0$	< 2.5 $\times 10^{-5}$	CL=90%	2128
$D_{sJ}(2457)^- K^+, D_{sJ}^- \rightarrow D_s^- \pi^0$	< 9.4 $\times 10^{-6}$	CL=90%	—
$D_{sJ}(2457)^- \pi^+, D_{sJ}^- \rightarrow D_s^- \pi^0$	< 4.0 $\times 10^{-6}$	CL=90%	—
$D_s^- D_s^+$	< 3.6 $\times 10^{-5}$	CL=90%	1759
$D_s^{*-} D_s^+$	< 1.3 $\times 10^{-4}$	CL=90%	1675
$D_s^{*-} D_s^{*+}$	< 2.4 $\times 10^{-4}$	CL=90%	1583
$D_{s0}^*(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0$	( 1.04 $\pm$ 0.17 ) $\times 10^{-3}$	S=1.1	1602
$D_{s0}(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^{*+} \gamma$	< 9.5 $\times 10^{-4}$	CL=90%	—
$D_{s0}(2317)^+ D^*(2010)^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0$	( 1.5 $\pm$ 0.6 ) $\times 10^{-3}$	—	1509
$D_{sJ}(2457)^+ D^-, D_{sJ}^{*+} \rightarrow D_s^+ \pi^0$	( 3.5 $\pm$ 1.1 ) $\times 10^{-3}$	—	—
$D_{sJ}(2457)^+ D^-, D_{sJ}^{*+} \rightarrow D_s^+ \gamma$	( 6.5 $\pm$ 1.7 ) $\times 10^{-4}$	—	—
$D_{sJ}(2457)^+ D^-, D_{sJ}^{*+} \rightarrow D_s^{*+} \gamma$	< 6.0 $\times 10^{-4}$	CL=90%	—
$D_{sJ}(2457)^+ D^-, D_{sJ}^{*+} \rightarrow D_s^+ \pi^+ \pi^-$	< 2.0 $\times 10^{-4}$	CL=90%	—
$D_{sJ}(2457)^+ D^-, D_{sJ}^{*+} \rightarrow D_s^+ \pi^0$	< 3.6 $\times 10^{-4}$	CL=90%	—
$D^*(2010)^- D_{sJ}(2457)^+$	( 9.3 $\pm$ 2.2 ) $\times 10^{-3}$	—	—

## Meson Summary Table

$D_{sJ}(2457)^+ D^*(2010), D_{sJ}^+ \rightarrow D_s^+ \gamma$	$(2.3 \pm 0.9) \times 10^{-3}$	—	$D^*(2010)^+ D^*(2010)^- \bar{D}^*(2007)^0 \omega$	$(8.0 \pm 0.6) \times 10^{-4}$	1711
$D^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+ + D^{*+} K^0$	$(2.8 \pm 0.7) \times 10^{-4}$	1444	$D^*(2010)^+ D^- D^*(2007)^0 \bar{D}^*(2007)^0$	$(3.6 \pm 1.1) \times 10^{-4}$	S=3.1 2180
$D^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+$	$(1.7 \pm 0.6) \times 10^{-4}$	1444	$D^- D^0 K^+$	$(6.1 \pm 1.5) \times 10^{-4}$	S=1.6 1790
$D^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0$	$(2.6 \pm 1.1) \times 10^{-4}$	1444	$D^- D^*(2007)^0 K^+$	$(1.07 \pm 0.11) \times 10^{-3}$	CL=90% 1715
$D^*(2010)^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+ + D^{*+} K^0$	$(5.0 \pm 1.4) \times 10^{-4}$	1336	$D^*(2010)^- D^0 K^+$	$(3.5 \pm 0.4) \times 10^{-3}$	1478
$D^*(2010)^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+$	$(3.3 \pm 1.1) \times 10^{-4}$	1336	$D^*(2010)^- D^*(2007)^0 K^+$	$(2.47 \pm 0.21) \times 10^{-3}$	1479
$D^*(2010)^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0$	$(5.0 \pm 1.7) \times 10^{-4}$	1336	$D^*(2010)^- D^*(2007)^0 K^+$	$(1.06 \pm 0.09) \%$	1366
$D^- D_{sJ}(2573)^+, D_{sJ}^+ \rightarrow D^0 K^+$	$(3.4 \pm 1.8) \times 10^{-5}$	1414	$D^- D^+ K^0$	$(7.5 \pm 1.7) \times 10^{-4}$	1568
$D^*(2010)^- D_{sJ}(2573)^+, D_{sJ}^+ \rightarrow D^0 K^+$	$< 2 \times 10^{-4}$	CL=90% 1304	$D^*(2010)^- D^+ K^0 + D^- D^*(2010)^+ K^0$	$(6.4 \pm 0.5) \times 10^{-3}$	1473
$D^- D_{sJ}(2700)^+, D_{sJ}^+ \rightarrow D^0 K^+$	$(7.1 \pm 1.2) \times 10^{-4}$	—	$D^*(2010)^- D^*(2010)^+ K^0$	$(8.1 \pm 0.7) \times 10^{-3}$	1360
$D^+ \pi^-$	$(7.4 \pm 1.3) \times 10^{-7}$	2306	$D^*(2010)^- D^*(2010)^+ K^0$	$(8.0 \pm 2.4) \times 10^{-4}$	1336
$D_s^+ \pi^-$	$(2.16 \pm 0.26) \times 10^{-5}$	2270	$\bar{D}^0 D^0 K^0$	$(2.7 \pm 1.1) \times 10^{-4}$	1574
$D_s^{*+} \pi^-$	$(2.1 \pm 0.4) \times 10^{-5}$	S=1.4 2215	$\bar{D}^0 D^*(2007)^0 K^0 + \bar{D}^*(2007)^0 D^0 K^0$	$(1.1 \pm 0.5) \times 10^{-3}$	1478
$D_s^+ \rho^-$	$< 2.4 \times 10^{-5}$	CL=90% 2197	$\bar{D}^*(2007)^0 D^0 K^0$	$(2.4 \pm 0.9) \times 10^{-3}$	1365
$D_s^{*+} \rho^-$	$(4.1 \pm 1.3) \times 10^{-5}$	2138	$(\bar{D} + \bar{D}^*)(D + D^*) K$	$(3.68 \pm 0.26) \%$	—
$D_s^+ a_0^-$	$< 1.9 \times 10^{-5}$	CL=90% —	<b>Charmonium modes</b>		
$D_s^+ a_0^-$	$< 3.6 \times 10^{-5}$	CL=90% —	$\eta_c K^0$	$(8.0 \pm 1.2) \times 10^{-4}$	1751
$D_s^+ a_1(1260)^-$	$< 2.1 \times 10^{-3}$	CL=90% 2080	$\eta_c K^*(892)^0$	$(6.3 \pm 0.9) \times 10^{-4}$	1646
$D_s^{*+} a_1(1260)^-$	$< 1.7 \times 10^{-3}$	CL=90% 2015	$\eta_c(2S) K^{*0}$	$< 3.9 \times 10^{-4}$	CL=90% 1157
$D_s^+ a_2^-$	$< 1.9 \times 10^{-4}$	CL=90% —	$h_c(1P) K^{*0}$	$< 4 \times 10^{-4}$	CL=90% 1253
$D_s^{*+} a_2^-$	$< 2.0 \times 10^{-4}$	CL=90% —	$J/\psi(1S) K^0$	$(8.73 \pm 0.32) \times 10^{-4}$	1683
$D_s^- K^+$	$(2.7 \pm 0.5) \times 10^{-5}$	S=2.7 2242	$J/\psi(1S) K^+ \pi^-$	$(1.15 \pm 0.05) \times 10^{-3}$	1652
$D_s^- K^+ K^+$	$(2.19 \pm 0.30) \times 10^{-5}$	2185	$J/\psi(1S) K^*(892)^0$	$(1.28 \pm 0.05) \times 10^{-3}$	1571
$D_s^- K^*(892)^+$	$(3.5 \pm 1.0) \times 10^{-5}$	2172	$J/\psi(1S) \eta K_S^0$	$(5.4 \pm 0.9) \times 10^{-5}$	1508
$D_s^{*-} K^*(892)^+$	$(3.2 \pm 1.5) \times 10^{-5}$	2112	$J/\psi(1S) \eta' K_S^0$	$< 2.5 \times 10^{-5}$	CL=90% 1271
$D_s^- \pi^+ K^0$	$(9.7 \pm 1.4) \times 10^{-5}$	2222	$J/\psi(1S) \phi K^0$	$(4.9 \pm 1.0) \times 10^{-5}$	S=1.3 1224
$D_s^{*-} \pi^+ K^0$	$< 1.10 \times 10^{-4}$	CL=90% 2164	$J/\psi(1S) \omega K^0$	$(2.3 \pm 0.4) \times 10^{-4}$	1386
$D_s^- K^+ \pi^+ \pi^-$	$(1.7 \pm 0.5) \times 10^{-4}$	2198	$X(3872) K^0, X \rightarrow J/\psi \omega$	$(6.0 \pm 3.2) \times 10^{-6}$	1140
$D_s^- \pi^+ K^*(892)^0$	$< 3.0 \times 10^{-3}$	CL=90% 2138	$X(3915), X \rightarrow J/\psi \omega$	$(2.1 \pm 0.9) \times 10^{-5}$	1102
$D_s^{*-} \pi^+ K^*(892)^0$	$< 1.6 \times 10^{-3}$	CL=90% 2076	$J/\psi(1S) K(1270)^0$	$(1.3 \pm 0.5) \times 10^{-3}$	1391
$\bar{D}^0 K^0$	$(5.2 \pm 0.7) \times 10^{-5}$	2280	$J/\psi(1S) \pi^0$	$(1.76 \pm 0.16) \times 10^{-5}$	S=1.1 1728
$\bar{D}^0 K^+ \pi^-$	$(8.8 \pm 1.7) \times 10^{-5}$	2261	$J/\psi(1S) \eta$	$(1.08 \pm 0.24) \times 10^{-5}$	S=1.5 1673
$\bar{D}^0 K^*(892)^0$	$(4.5 \pm 0.6) \times 10^{-5}$	2213	$J/\psi(1S) \pi^+ \pi^-$	$(4.03 \pm 0.18) \times 10^{-5}$	1716
$\bar{D}^0 K^*(1410)^0$	$< 6.7 \times 10^{-5}$	CL=90% 2062	$J/\psi(1S) \pi^+ \pi^- \text{ nonresonant}$	$< 1.2 \times 10^{-5}$	CL=90% 1716
$\bar{D}^0 K_2^*(1430)^0$	$(7 \pm 7) \times 10^{-6}$	2057	$J/\psi(1S) f_0(500), f_0 \rightarrow \pi \pi$	$(8.1 \pm 1.1) \times 10^{-6}$	—
$D_0^*(2400)^-, D_0^{*-} \rightarrow \bar{D}^0 \pi^-$	$(1.9 \pm 0.9) \times 10^{-5}$	—	$J/\psi(1S) f_2$	$(3.3 \pm 0.5) \times 10^{-6}$	S=1.6 —
$D_2^*(2460)^- K^+, D_2^{*-} \rightarrow \bar{D}^0 \pi^-$	$(2.03 \pm 0.35) \times 10^{-5}$	2029	$J/\psi(1S) \rho^0$	$(2.54 \pm 0.14) \times 10^{-5}$	1612
$D_3^*(2760)^- K^+, D_3^{*-} \rightarrow \bar{D}^0 \pi^-$	$< 1.0 \times 10^{-6}$	CL=90% —	$J/\psi(1S) f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 1.1 \times 10^{-6}$	CL=90% —
$\bar{D}^0 K^+ \pi^- \text{ non-resonant}$	$< 3.7 \times 10^{-5}$	CL=90% —	$J/\psi(1S) \rho(1450)^0, \rho^0 \rightarrow \pi^+ \pi^-$	$(3.0 \pm 1.6) \times 10^{-6}$	—
$[K^+ K^-]_D K^*(892)^0$	$(4.7 \pm 0.9) \times 10^{-5}$	—	$J/\psi \pi \pi$	$(2.0 \pm 1.3) \times 10^{-6}$	—
$[\pi^+ \pi^-]_D K^*(892)^0$	$(5.5 \pm 1.4) \times 10^{-5}$	—	$J/\psi(1S) \omega$	$(1.8 \pm 0.7) \times 10^{-5}$	1609
$\bar{D}^0 \pi^0$	$(2.63 \pm 0.14) \times 10^{-4}$	2308	$J/\psi(1S) K^+ K^-$	$(2.6 \pm 0.4) \times 10^{-6}$	1533
$\bar{D}^0 \rho^0$	$(3.21 \pm 0.21) \times 10^{-4}$	2237	$J/\psi(1S) a_0(980), a_0 \rightarrow K^+ K^-$	$(4.7 \pm 3.4) \times 10^{-7}$	—
$\bar{D}^0 f_2$	$(1.56 \pm 0.21) \times 10^{-4}$	—	$J/\psi(1S) \phi$	$< 1.9 \times 10^{-7}$	CL=90% 1520
$\bar{D}^0 \eta$	$(2.36 \pm 0.32) \times 10^{-4}$	S=2.5 2274	$J/\psi(1S) \eta'(958)$	$(7.6 \pm 2.4) \times 10^{-6}$	1546
$\bar{D}^0 \eta'$	$(1.38 \pm 0.16) \times 10^{-4}$	S=1.3 2198	$J/\psi(1S) K^0 \pi^+ \pi^-$	$(4.4 \pm 0.4) \times 10^{-4}$	1611
$\bar{D}^0 \omega$	$(2.54 \pm 0.16) \times 10^{-4}$	2235	$J/\psi(1S) K^0 K^+ \pi^-$	$< 2.1 \times 10^{-5}$	CL=90% 1467
$D^0 \phi$	$< 1.16 \times 10^{-5}$	CL=90% 2183	$J/\psi(1S) K^0 K^+ K^-$	$(2.5 \pm 0.7) \times 10^{-5}$	S=1.8 1249
$D^0 K^+ \pi^-$	$(5.3 \pm 3.2) \times 10^{-6}$	2261	$J/\psi(1S) K^0 \rho^0$	$(5.4 \pm 3.0) \times 10^{-4}$	1390
$D^0 K^*(892)^0$	$< 1.1 \times 10^{-5}$	CL=90% 2213	$J/\psi(1S) K^*(892)^+ \pi^-$	$(8 \pm 4) \times 10^{-4}$	1514
$\bar{D}^{*0} \gamma$	$< 2.5 \times 10^{-5}$	CL=90% 2258	$J/\psi(1S) \pi^+ \pi^- \pi^+$	$(1.45 \pm 0.13) \times 10^{-5}$	1670
$\bar{D}^*(2007)^0 \pi^0$	$(2.2 \pm 0.6) \times 10^{-4}$	S=2.6 2256	$J/\psi(1S) f_1(1285)$	$(8.4 \pm 2.1) \times 10^{-6}$	1385
$\bar{D}^*(2007)^0 \rho^0$	$< 5.1 \times 10^{-4}$	CL=90% 2182	$J/\psi(1S) K^*(892)^0 \pi^+ \pi^-$	$(6.6 \pm 2.2) \times 10^{-4}$	1447
$\bar{D}^*(2007)^0 \eta$	$(2.3 \pm 0.6) \times 10^{-4}$	S=2.8 2220	$X(3872)^- K^+, X(3872)^- \rightarrow [zzz]$	$< 5 \times 10^{-4}$	CL=90% —
$\bar{D}^*(2007)^0 \eta'$	$(1.40 \pm 0.22) \times 10^{-4}$	2141	$X(3872)^- K^+, X(3872)^- \rightarrow [zzz]$	$< 4.2 \times 10^{-6}$	CL=90% —
$\bar{D}^*(2007)^0 \pi^+ \pi^-$	$(6.2 \pm 2.2) \times 10^{-4}$	2249	$J/\psi(1S) \pi^- \pi^0$	$(4.3 \pm 1.3) \times 10^{-6}$	1140
$\bar{D}^*(2007)^0 K^0$	$(3.6 \pm 1.2) \times 10^{-5}$	2227	$X(3872) K^0, X \rightarrow J/\psi \pi^+ \pi^-$	$< 2.4 \times 10^{-6}$	CL=90% 1140
$\bar{D}^*(2007)^0 K^*(892)^0$	$< 6.9 \times 10^{-5}$	CL=90% 2157	$X(3872) K^0, X \rightarrow J/\psi \gamma$	$< 2.8 \times 10^{-6}$	CL=90% 940
$D^*(2007)^0 K^*(892)^0$	$< 4.0 \times 10^{-5}$	CL=90% 2157	$X(3872) K^*(892)^0, X \rightarrow J/\psi \gamma$	$< 6.62 \times 10^{-6}$	CL=90% 1140
$D^*(2007)^0 \pi^+ \pi^- \pi^-$	$(2.7 \pm 0.5) \times 10^{-3}$	2219	$X(3872) K^*(892)^0, X \rightarrow \psi(2S) \gamma$	$< 4.4 \times 10^{-6}$	CL=90% 940
			$\psi(2S) \gamma$	$(1.7 \pm 0.8) \times 10^{-4}$	1140
			$X(3872) K^0, X \rightarrow D^0 \bar{D}^0 \pi^0$	$(1.2 \pm 0.4) \times 10^{-4}$	1140
			$X(3872) K^+ \pi^-, X \rightarrow J/\psi \pi^+ \pi^-$	$(7.9 \pm 1.4) \times 10^{-6}$	—

# Meson Summary Table

$X(3872)K^*(982)^0, X \rightarrow J/\psi \pi^+ \pi^-$	$(4.0 \pm 1.5) \times 10^{-6}$	—	$K^*(1410)^+ \pi^-, K^{*+} \rightarrow K^0 \pi^+$	$< 3.8 \times 10^{-6}$	CL=90%	—
$X(4430)^\pm K^\mp, X^\pm \rightarrow \psi(2S)\pi^\pm$	$(6.0 \pm_{-2.4}^{3.0}) \times 10^{-5}$	583	$f_0(980)K^0, f_0 \rightarrow \pi^+ \pi^-$	$(7.0 \pm 0.9) \times 10^{-6}$		2522
$X(4430)^\pm K^\mp, X^\pm \rightarrow J/\psi \pi^\pm$	$(5.4 \pm_{-1.2}^{4.0}) \times 10^{-6}$	583	$f_2(1270)K^0$	$(2.7 \pm_{-1.2}^{1.3}) \times 10^{-6}$		2459
$X(3900)^\pm K^\mp, X^\pm \rightarrow J/\psi \pi^\pm$	$< 9 \times 10^{-7}$	—	$f_x(1300)K^0, f_x \rightarrow \pi^+ \pi^-$	$(1.8 \pm 0.7) \times 10^{-6}$		—
$X(4200)^\pm K^\mp, X^\pm \rightarrow J/\psi \pi^\pm$	$(2.2 \pm_{-0.8}^{1.3}) \times 10^{-5}$	—	$K^*(892)^0 \pi^0$	$(3.3 \pm 0.6) \times 10^{-6}$		2563
$J/\psi(1S)p\bar{p}$	$< 5.2 \times 10^{-7}$	CL=90% 862	$K_2^*(1430)^+ \pi^-$	$< 6 \times 10^{-6}$	CL=90%	2445
$J/\psi(1S)\gamma$	$< 1.5 \times 10^{-6}$	CL=90% 1732	$K^*(1680)^+ \pi^-$	$< 1.0 \times 10^{-5}$	CL=90%	2358
$J/\psi(1S)\bar{D}^0$	$< 1.3 \times 10^{-5}$	CL=90% 877	$K^+ \pi^- \pi^+ \pi^-$	$[ddaa] < 2.3 \times 10^{-4}$	CL=90%	2600
$\psi(2S)\pi^0$	$(1.17 \pm 0.19) \times 10^{-5}$	1348	$\rho^0 K^+ \pi^-$	$(2.8 \pm 0.7) \times 10^{-6}$		2543
$\psi(2S)K^0$	$(5.8 \pm 0.5) \times 10^{-4}$	1283	$f_0(980)K^+ \pi^-, f_0 \rightarrow \pi\pi$	$(1.4 \pm_{-0.6}^{0.5}) \times 10^{-6}$		2506
$\psi(3770)K^0, \psi \rightarrow \bar{D}^0 D^0$	$< 1.23 \times 10^{-4}$	CL=90% 1217	$K^+ \pi^- \pi^+ \pi^-$ nonresonant	$< 2.1 \times 10^{-6}$	CL=90%	2600
$\psi(3770)K^0, \psi \rightarrow D^- D^+$	$< 1.88 \times 10^{-4}$	CL=90% 1217	$K^*(892)^0 \pi^+ \pi^-$	$(5.5 \pm 0.5) \times 10^{-5}$		2557
$\psi(2S)\pi^+ \pi^-$	$(2.3 \pm 0.4) \times 10^{-5}$	1331	$K^*(892)^0 \rho^0$	$(3.9 \pm 1.3) \times 10^{-6}$	S=1.9	2504
$\psi(2S)K^+ \pi^-$	$(5.8 \pm 0.4) \times 10^{-4}$	1239	$K^*(892)^0 f_0(980), f_0 \rightarrow \pi\pi$	$(3.9 \pm_{-1.8}^{2.1}) \times 10^{-6}$	S=3.9	2466
$\psi(2S)K^*(892)^0$	$(5.9 \pm 0.4) \times 10^{-4}$	1116	$K_1(1270)^+ \pi^-$	$< 3.0 \times 10^{-5}$	CL=90%	2484
$\chi_{c0}K^0$	$(1.47 \pm 0.27) \times 10^{-4}$	1477	$K_1(1400)^+ \pi^-$	$< 2.7 \times 10^{-5}$	CL=90%	2451
$\chi_{c0}K^*(892)^0$	$(1.7 \pm 0.4) \times 10^{-4}$	1342	$a_1(1260)^- K^+$	$[ddaa] (1.6 \pm 0.4) \times 10^{-5}$		2471
$\chi_{c2}K^0$	$< 1.5 \times 10^{-5}$	CL=90% 1379	$K^*(892)^+ \rho^-$	$(1.03 \pm 0.26) \times 10^{-5}$		2504
$\chi_{c2}K^*(892)^0$	$(4.9 \pm 1.2) \times 10^{-5}$	S=1.1 1228	$K_0^*(1430)^+ \rho^-$	$(2.8 \pm 1.2) \times 10^{-5}$		—
$\chi_{c1}\pi^0$	$(1.12 \pm 0.28) \times 10^{-5}$	1468	$K_1^*(1400)^0 \rho^0$	$< 3.0 \times 10^{-3}$	CL=90%	2388
$\chi_{c1}K^0$	$(3.93 \pm 0.27) \times 10^{-4}$	1411	$K_0^*(1430)^0 \rho^0$	$(2.7 \pm 0.6) \times 10^{-5}$		2381
$\chi_{c1}K^- \pi^+$	$(3.8 \pm 0.4) \times 10^{-4}$	1371	$K_0^*(1430)^0 f_0(980), f_0 \rightarrow \pi\pi$	$(2.7 \pm 0.9) \times 10^{-6}$		—
$\chi_{c1}K^*(892)^0$	$(2.39 \pm 0.19) \times 10^{-4}$	S=1.2 1265	$K_2^*(1430)^0 f_0(980), f_0 \rightarrow \pi\pi$	$(8.6 \pm 2.0) \times 10^{-6}$		—
$X(4051)^+ K^-, X^+ \rightarrow \chi_{c1}\pi^+$	$(3.0 \pm_{-1.8}^{4.0}) \times 10^{-5}$	—	$K^+ K^-$	$(1.3 \pm 0.5) \times 10^{-7}$		2593
$X(4248)^+ K^-, X^+ \rightarrow \chi_{c1}\pi^+$	$(4.0 \pm_{-1.0}^{20.0}) \times 10^{-5}$	—	$K^0 \bar{K}^0$	$(1.21 \pm 0.16) \times 10^{-6}$		2592
<b>K or K* modes</b>			$K^0 K^- \pi^+$	$(6.5 \pm 0.8) \times 10^{-6}$		2578
$K^+ \pi^-$	$(1.96 \pm 0.05) \times 10^{-5}$	2615	$K^*(892)^\pm K^\mp$	$< 4 \times 10^{-7}$	CL=90%	2540
$K^0 \pi^0$	$(9.9 \pm 0.5) \times 10^{-6}$	2615	$\bar{K}^{*0} K^0 + K^{*0} \bar{K}^0$	$< 9.6 \times 10^{-7}$	CL=90%	—
$\eta' K^0$	$(6.6 \pm 0.4) \times 10^{-5}$	S=1.4 2528	$K^+ K^- \pi^0$	$(2.2 \pm 0.6) \times 10^{-6}$		2579
$\eta' K^*(892)^0$	$(2.8 \pm 0.6) \times 10^{-6}$	2472	$K^0 K_S^0 \pi^0$	$< 9 \times 10^{-7}$	CL=90%	2578
$\eta' K_0^*(1430)^0$	$(6.3 \pm 1.6) \times 10^{-6}$	2346	$K_S^0 K_S^0 \eta$	$< 1.0 \times 10^{-6}$	CL=90%	2515
$\eta' K_2^*(1430)^0$	$(1.37 \pm 0.32) \times 10^{-5}$	2346	$K_S^0 K_S^0 \eta'$	$< 2.0 \times 10^{-6}$	CL=90%	2452
$\eta K^0$	$(1.23 \pm_{-0.24}^{0.27}) \times 10^{-6}$	2587	$K^0 K^+ K^-$	$(2.49 \pm 0.31) \times 10^{-5}$	S=3.0	2522
$\eta K^*(892)^0$	$(1.59 \pm 0.10) \times 10^{-5}$	2534	$K^0 \phi$	$(7.3 \pm 0.7) \times 10^{-6}$		2516
$\eta K_0^*(1430)^0$	$(1.10 \pm 0.22) \times 10^{-5}$	2415	$f_0(980)K^0, f_0 \rightarrow K^+ K^-$	$(7.0 \pm_{-3.0}^{3.5}) \times 10^{-6}$		—
$\eta K_2^*(1430)^0$	$(9.6 \pm 2.1) \times 10^{-6}$	2414	$f_0(1500)K^0$	$(1.3 \pm_{-0.5}^{0.7}) \times 10^{-5}$		2398
$\omega K^0$	$(4.8 \pm 0.4) \times 10^{-6}$	2557	$f_2'(1525)^0 K^0$	$(3 \pm_{-4}^5) \times 10^{-7}$		—
$a_0(980)^0 K^0, a_0^0 \rightarrow \eta \pi^0$	$< 7.8 \times 10^{-6}$	CL=90% —	$f_0(1710)K^0, f_0 \rightarrow K^+ K^-$	$(4.4 \pm 0.9) \times 10^{-6}$		—
$b_1^0 K^0, b_1^0 \rightarrow \omega \pi^0$	$< 7.8 \times 10^{-6}$	CL=90% —	$K^0 K^+ K^-$ nonresonant	$(3.3 \pm 1.0) \times 10^{-5}$		2522
$a_0(980)^\pm K^\mp, a_0^\pm \rightarrow \eta \pi^\pm$	$< 1.9 \times 10^{-6}$	CL=90% —	$K_S^0 K_S^0 K_S^0$	$(6.0 \pm 0.5) \times 10^{-6}$	S=1.1	2521
$b_1^- K^+, b_1^- \rightarrow \omega \pi^-$	$(7.4 \pm 1.4) \times 10^{-6}$	—	$f_0(980)K^0, f_0 \rightarrow K_S^0 K_S^0$	$(2.7 \pm 1.8) \times 10^{-6}$		—
$b_1^0 K^{*0}, b_1^0 \rightarrow \omega \pi^0$	$< 8.0 \times 10^{-6}$	CL=90% —	$f_0(1710)K^0, f_0 \rightarrow K_S^0 K_S^0$	$(5.0 \pm_{-2.6}^{5.0}) \times 10^{-7}$		—
$b_1^- K^{*+}, b_1^- \rightarrow \omega \pi^-$	$< 5.0 \times 10^{-6}$	CL=90% —	$f_0(2010)K^0, f_0 \rightarrow K_S^0 K_S^0$	$(5 \pm 6) \times 10^{-7}$		—
$a_0(1450)^\pm K^\mp, a_0^\pm \rightarrow \eta \pi^\pm$	$< 3.1 \times 10^{-6}$	CL=90% —	$K_S^0 K_S^0 K_S^0$ nonresonant	$(1.33 \pm 0.31) \times 10^{-5}$		2521
$K_S^0 X^0$ (Familon)	$< 5.3 \times 10^{-5}$	CL=90% —	$K_S^0 K_S^0 K_L^0$	$< 1.6 \times 10^{-5}$	CL=90%	2521
$\omega K^*(892)^0$	$(2.0 \pm 0.5) \times 10^{-6}$	2503	$K^*(892)^0 K^+ K^-$	$(2.75 \pm 0.26) \times 10^{-5}$		2467
$\omega(K\pi)_0^0$	$(1.84 \pm 0.25) \times 10^{-5}$	—	$K^*(892)^0 \phi$	$(1.00 \pm 0.05) \times 10^{-5}$		2460
$\omega K_0^*(1430)^0$	$(1.60 \pm 0.34) \times 10^{-5}$	2380	$K^+ K^- \pi^+ \pi^-$ nonresonant	$< 7.17 \times 10^{-5}$	CL=90%	2559
$\omega K_2^*(1430)^0$	$(1.01 \pm 0.23) \times 10^{-5}$	2380	$K^*(892)^0 K^- \pi^+$	$(4.5 \pm 1.3) \times 10^{-6}$		2524
$\omega K^+ \pi^-$ nonresonant	$(5.1 \pm 1.0) \times 10^{-6}$	2542	$K^*(892)^0 \bar{K}^*(892)^0$	$(8 \pm 5) \times 10^{-7}$	S=2.2	2485
$K^+ \pi^- \pi^0$	$(3.78 \pm 0.32) \times 10^{-5}$	2609	$K^+ K^+ \pi^- \pi^-$ nonresonant	$< 6.0 \times 10^{-6}$	CL=90%	2559
$K^+ \rho^-$	$(7.0 \pm 0.9) \times 10^{-6}$	2559	$K^*(892)^0 K^+ \pi^-$	$< 2.2 \times 10^{-6}$	CL=90%	2524
$K^+ \rho(1450)^-$	$(2.4 \pm 1.2) \times 10^{-6}$	—	$K^*(892)^0 K^*(892)^0$	$< 2 \times 10^{-7}$	CL=90%	2485
$K^+ \rho(1700)^-$	$(6 \pm 7) \times 10^{-7}$	—	$K^*(892)^+ K^*(892)^-$	$< 2.0 \times 10^{-6}$	CL=90%	2485
$(K^+ \pi^- \pi^0)$ non-resonant	$(2.8 \pm 0.6) \times 10^{-6}$	—	$K_1(1400)^0 \phi$	$< 5.0 \times 10^{-3}$	CL=90%	2339
$(K\pi)_0^{*+} \pi^-, (K\pi)_0^{*+} \rightarrow$	$(3.4 \pm 0.5) \times 10^{-5}$	—	$\phi(K\pi)_0^{*0}$	$(4.3 \pm 0.4) \times 10^{-6}$		—
$(K\pi)_0^{*0} \pi^0, (K\pi)_0^{*0} \rightarrow$	$(8.6 \pm 1.7) \times 10^{-6}$	—	$\phi(K\pi)_0^{*0} (1.60 < m_{K\pi} < 2.15]_{eeaa}]$	$< 1.7 \times 10^{-6}$	CL=90%	—
$K^+ \pi^-$	$(4.0 \pm 0.5) \times 10^{-6}$	CL=90% 2445	$K_0^*(1430)^0 K^- \pi^+$	$< 3.18 \times 10^{-5}$	CL=90%	2403
$K_2^*(1430)^0 \pi^0$	$< 7.5 \times 10^{-6}$	CL=90% 2358	$K_0^*(1430)^0 \bar{K}^*(892)^0$	$< 3.3 \times 10^{-6}$	CL=90%	2360
$K^*(1680)^0 \pi^0$	$[ccaa] (6.1 \pm 1.6) \times 10^{-6}$	—	$K_0^*(1430)^0 \bar{K}_0^*(1430)^0$	$< 8.4 \times 10^{-6}$	CL=90%	2222
$K_x^{*0} \pi^0$	$(5.20 \pm 0.24) \times 10^{-5}$	S=1.3 2609	$K_0^*(1430)^0 \phi$	$(3.9 \pm 0.8) \times 10^{-6}$		2333
$K^0 \pi^+ \pi^-$	$(1.47 \pm_{-0.26}^{0.40}) \times 10^{-5}$	S=2.1 —	$K_0^*(1430)^0 K^*(892)^0$	$< 1.7 \times 10^{-6}$	CL=90%	2360
$K^0 \pi^+ \pi^-$ non-resonant	$(4.7 \pm 0.6) \times 10^{-6}$	2558	$K_0^*(1430)^0 K_0^*(1430)^0$	$< 4.7 \times 10^{-6}$	CL=90%	2222
$K^0 \rho^0$	$(8.4 \pm 0.8) \times 10^{-6}$	2563	$K^*(1680)^0 \phi$	$< 3.5 \times 10^{-6}$	CL=90%	2238
$K^*(892)^+ \pi^-$	$(3.3 \pm 0.7) \times 10^{-5}$	S=2.0 —	$K^*(1780)^0 \phi$	$< 2.7 \times 10^{-6}$	CL=90%	—
$K_0^*(1430)^+ \pi^-$	$(5.1 \pm 1.6) \times 10^{-6}$	—	$K^*(2045)^0 \phi$	$< 1.53 \times 10^{-5}$	CL=90%	—
$K_x^{*+} \pi^-$	$[ccaa] (5.1 \pm 1.6) \times 10^{-6}$	—	$K_2^*(1430)^0 \rho^0$	$< 1.1 \times 10^{-3}$	CL=90%	2381
			$K_2^*(1430)^0 \phi$	$(6.8 \pm 0.9) \times 10^{-6}$	S=1.2	2333
			$K^0 \phi \phi$	$(4.5 \pm 0.9) \times 10^{-6}$		2305
			$\eta' \eta' K^0$	$< 3.1 \times 10^{-5}$	CL=90%	2337



## Meson Summary Table

$\eta K^0 \gamma$	$(7.6 \pm 1.8) \times 10^{-6}$	2587	$\pi^+ \pi^+ \pi^+ \pi^- \pi^- \pi^-$	$< 3.0 \times 10^{-3}$	CL=90%	2592
$\eta' K^0 \gamma$	$< 6.4 \times 10^{-6}$	2528	$a_1(1260)^+ a_1(1260)^-, a_1^+ \rightarrow$	$(1.18 \pm 0.31) \times 10^{-5}$		2336
$K^0 \phi \gamma$	$(2.7 \pm 0.7) \times 10^{-6}$	2516	$2\pi^+ \pi^-, a_1^- \rightarrow 2\pi^- \pi^+$			
$K^+ \pi^- \gamma$	$(4.6 \pm 1.4) \times 10^{-6}$	2615	$\pi^+ \pi^+ \pi^+ \pi^- \pi^- \pi^0$	$< 1.1 \%$	CL=90%	2572
$K^*(892)^0 \gamma$	$(4.33 \pm 0.15) \times 10^{-5}$	2565	<b>Baryon modes</b>			
$K^*(1410) \gamma$	$< 1.3 \times 10^{-4}$	2451	$p\bar{p}$	$(1.5 \pm 0.7) \times 10^{-8}$		2467
$K^+ \pi^- \gamma$ nonresonant	$< 2.6 \times 10^{-6}$	2615	$p\bar{p} \pi^+ \pi^-$	$< 2.5 \times 10^{-4}$	CL=90%	2406
$K^*(892)^0 X(214), X \rightarrow$	$[f\bar{f}aa] < 2.26 \times 10^{-8}$	CL=90%	$p\bar{p} K^0$	$(2.66 \pm 0.32) \times 10^{-6}$		2347
$\mu^+ \mu^-$			$\Theta(1540)^+ \bar{p}, \Theta^+ \rightarrow p K_S^0$	$< 5 \times 10^{-8}$	CL=90%	2318
$K^0 \pi^+ \pi^- \gamma$	$(1.95 \pm 0.22) \times 10^{-5}$	2609	$f_J(2220) K^0, f_J \rightarrow p\bar{p}$	$< 4.5 \times 10^{-7}$	CL=90%	2135
$K^+ \pi^- \pi^0 \gamma$	$(4.1 \pm 0.4) \times 10^{-5}$	2609	$p\bar{p} K^*(892)^0$	$(1.24 \pm 0.28) \times 10^{-6}$		2216
$K_1(1270)^0 \gamma$	$< 5.8 \times 10^{-5}$	2486	$f_J(2220) K_0^*, f_J \rightarrow p\bar{p}$	$< 1.5 \times 10^{-7}$	CL=90%	—
$K_1(1400)^0 \gamma$	$< 1.2 \times 10^{-5}$	2454	$p\bar{p} \pi^-$	$(3.14 \pm 0.29) \times 10^{-6}$		2401
$K_2^*(1430)^0 \gamma$	$(1.24 \pm 0.24) \times 10^{-5}$	2447	$p\bar{p} \pi^- \gamma$	$< 6.5 \times 10^{-7}$	CL=90%	2401
$K^*(1680)^0 \gamma$	$< 2.0 \times 10^{-3}$	2361	$p\bar{p} \Sigma(1385)^-$	$< 2.6 \times 10^{-7}$	CL=90%	2363
$K_3^*(1780)^0 \gamma$	$< 8.3 \times 10^{-5}$	2341	$\Delta^0 \bar{\Lambda}$	$< 9.3 \times 10^{-7}$	CL=90%	2364
$K_4^*(2045)^0 \gamma$	$< 4.3 \times 10^{-3}$	2244	$p\bar{p} \Lambda K^-$	$< 8.2 \times 10^{-7}$	CL=90%	2308
<b>Light unflavored meson modes</b>			$p\bar{p} D^-$	$(2.5 \pm 0.4) \times 10^{-5}$		1765
$\rho^0 \gamma$	$(8.6 \pm 1.5) \times 10^{-7}$	2583	$p\bar{p} D^{*-}$	$(3.4 \pm 0.8) \times 10^{-5}$		1685
$\rho^0 X(214), X \rightarrow \mu^+ \mu^-$	$[f\bar{f}aa] < 1.73 \times 10^{-8}$	CL=90%	$p\bar{p} \Sigma^0 \pi^-$	$< 3.8 \times 10^{-6}$	CL=90%	2383
$\omega \gamma$	$(4.4 \pm 1.8) \times 10^{-7}$	2582	$\bar{\Lambda} \Lambda$	$< 3.2 \times 10^{-7}$	CL=90%	2392
$\phi \gamma$	$< 8.5 \times 10^{-7}$	CL=90%	$\bar{\Lambda} \Lambda K^0$	$(4.8 \pm 1.0) \times 10^{-6}$		2250
$\pi^+ \pi^-$	$(5.12 \pm 0.19) \times 10^{-6}$	2636	$\bar{\Lambda} \Lambda K^{*0}$	$(2.5 \pm 0.9) \times 10^{-6}$		2098
$\pi^0 \pi^0$	$(1.91 \pm 0.22) \times 10^{-6}$	2636	$\bar{\Lambda} \Lambda D^0$	$(1.00 \pm 0.30) \times 10^{-5}$		1661
$\eta \pi^0$	$(4.1 \pm 1.7) \times 10^{-7}$	2610	$D^0 \Sigma^0 \bar{\Lambda} + \text{c.c.}$	$< 3.1 \times 10^{-5}$	CL=90%	1611
$\eta \eta$	$< 1.0 \times 10^{-6}$	CL=90%	$\Delta^0 \bar{\Delta}^0$	$< 1.5 \times 10^{-3}$	CL=90%	2335
$\eta' \pi^0$	$(1.2 \pm 0.6) \times 10^{-6}$	S=1.7 2551	$\Delta^{++} \bar{\Delta}^{--}$	$< 1.1 \times 10^{-4}$	CL=90%	2335
$\eta' \eta'$	$< 1.7 \times 10^{-6}$	CL=90%	$\bar{D}^0 p\bar{p}$	$(1.04 \pm 0.07) \times 10^{-4}$		1863
$\eta' \eta$	$< 1.2 \times 10^{-6}$	CL=90%	$D_s^- \bar{\Lambda} p$	$(2.8 \pm 0.9) \times 10^{-5}$		1710
$\eta' \rho^0$	$< 1.3 \times 10^{-6}$	CL=90%	$\bar{D}^*(2007)^0 p\bar{p}$	$(9.9 \pm 1.1) \times 10^{-5}$		1788
$\eta' f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 9 \times 10^{-7}$	CL=90%	$D^*(2010)^- p\bar{n}$	$(1.4 \pm 0.4) \times 10^{-3}$		1785
$\eta \rho^0$	$< 1.5 \times 10^{-6}$	CL=90%	$D^- p\bar{p} \pi^+$	$(3.32 \pm 0.31) \times 10^{-4}$		1786
$\eta f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 4 \times 10^{-7}$	CL=90%	$\bar{D}^*(2010)^- p\bar{p} \pi^+$	$(4.7 \pm 0.5) \times 10^{-4}$	S=1.2	1708
$\omega \eta$	$(9.4 \pm 4.0) \times 10^{-7}$	2552	$\bar{D}^0 p\bar{p} \pi^+ \pi^-$	$(3.0 \pm 0.5) \times 10^{-4}$		1708
$\omega \eta'$	$(1.0 \pm 0.5) \times 10^{-6}$	2491	$\bar{D}^{*0} p\bar{p} \pi^+ \pi^-$	$(1.9 \pm 0.5) \times 10^{-4}$		1623
$\omega \rho^0$	$< 1.6 \times 10^{-6}$	CL=90%	$\Theta_c \bar{p} \pi^+, \Theta_c \rightarrow D^- p$	$< 9 \times 10^{-6}$	CL=90%	—
$\omega f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 1.5 \times 10^{-6}$	CL=90%	$\Theta_c \bar{p} \pi^+, \Theta_c \rightarrow D^{*-} p$	$< 1.4 \times 10^{-5}$	CL=90%	—
$\omega \omega$	$(1.2 \pm 0.4) \times 10^{-6}$	2521	$\Sigma_c^- \Delta^{++}$	$< 8 \times 10^{-4}$	CL=90%	1839
$\phi \pi^0$	$< 1.5 \times 10^{-7}$	CL=90%	$\bar{\Lambda}_c^- p \pi^+ \pi^-$	$(1.01 \pm 0.14) \times 10^{-3}$	S=1.3	1934
$\phi \eta$	$< 5 \times 10^{-7}$	CL=90%	$\bar{\Lambda}_c^- p$	$(1.52 \pm 0.18) \times 10^{-5}$		2021
$\phi \eta'$	$< 5 \times 10^{-7}$	CL=90%	$\bar{\Lambda}_c^- p \pi^0$	$(1.53 \pm 0.18) \times 10^{-4}$		1982
$\phi \rho^0$	$< 3.3 \times 10^{-7}$	CL=90%	$\Sigma_c^-(2455)^- p$	$< 2.4 \times 10^{-5}$		—
$\phi f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 3.8 \times 10^{-7}$	CL=90%	$\bar{\Lambda}_c^- p \pi^+ \pi^- \pi^0$	$< 5.07 \times 10^{-3}$	CL=90%	1882
$\phi \omega$	$< 7 \times 10^{-7}$	CL=90%	$\bar{\Lambda}_c^- p \pi^+ \pi^- \pi^+ \pi^-$	$< 2.74 \times 10^{-3}$	CL=90%	1821
$\phi \phi$	$< 2.8 \times 10^{-8}$	CL=90%	$\bar{\Lambda}_c^- p \pi^+ \pi^- (\text{nonresonant})$	$(5.4 \pm 1.0) \times 10^{-4}$	S=1.3	1934
$a_0(980)^\pm \pi^\mp, a_0^\pm \rightarrow \eta \pi^\pm$	$< 3.1 \times 10^{-6}$	CL=90%	$\Sigma_c^-(2520)^- p \pi^+$	$(1.01 \pm 0.18) \times 10^{-4}$		1860
$a_0(1450)^\pm \pi^\mp, a_0^\pm \rightarrow \eta \pi^\pm$	$< 2.3 \times 10^{-6}$	CL=90%	$\Sigma_c^-(2520)^0 p \pi^-$	$< 3.1 \times 10^{-5}$	CL=90%	1860
$\pi^+ \pi^- \pi^0$	$< 7.2 \times 10^{-4}$	CL=90%	$\Sigma_c^-(2455)^0 p \pi^-$	$(1.07 \pm 0.16) \times 10^{-4}$		1895
$\rho^0 \pi^0$	$(2.0 \pm 0.5) \times 10^{-6}$	2581	$\Sigma_c^-(2455)^0 N^0, N^0 \rightarrow$	$(6.3 \pm 1.6) \times 10^{-5}$		—
$\rho^\mp \pi^\pm$	$[hh] (2.30 \pm 0.23) \times 10^{-5}$	2581	$\bar{\Sigma}_c^-(2455)^- p \pi^+$	$(1.81 \pm 0.24) \times 10^{-4}$		1895
$\pi^+ \pi^- \pi^+ \pi^-$	$< 1.12 \times 10^{-5}$	CL=90%	$\Lambda_c^- p K^+ \pi^-$	$(3.4 \pm 0.7) \times 10^{-5}$		—
$\rho^0 \pi^+ \pi^-$	$< 8.8 \times 10^{-6}$	CL=90%	$\bar{\Sigma}_c^-(2455)^- p K^+, \bar{\Sigma}_c^{--} \rightarrow$	$(8.7 \pm 2.5) \times 10^{-6}$		1754
$\rho^0 \rho^0$	$(9.6 \pm 1.5) \times 10^{-7}$	2523	$\bar{\Lambda}_c^- \pi^-$			
$f_0(980) \pi^+ \pi^-, f_0 \rightarrow$	$< 3.0 \times 10^{-6}$	CL=90%	$\Lambda_c^- p K^*(892)^0$	$< 2.42 \times 10^{-5}$	CL=90%	—
$\pi^+ \pi^-$			$\Lambda_c^- p K^+ K^-$	$(2.0 \pm 0.4) \times 10^{-5}$		—
$\rho^0 f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$(7.8 \pm 2.5) \times 10^{-7}$	2486	$\Lambda_c^- p \phi$	$< 9 \times 10^{-6}$	CL=90%	—
$f_0(980) f_0(980), f_0 \rightarrow$	$< 1.9 \times 10^{-7}$	CL=90%	$\Lambda_c^- p \bar{p} p$	$< 2.8 \times 10^{-6}$		—
$\pi^+ \pi^-, f_0 \rightarrow \pi^+ \pi^-$			$\bar{\Lambda}_c^- \Lambda K^+$	$(4.8 \pm 1.1) \times 10^{-5}$		1767
$f_0(980) f_0(980), f_0 \rightarrow \pi^+ \pi^-,$	$< 2.3 \times 10^{-7}$	CL=90%	$\bar{\Lambda}_c^- \Lambda_c^+$	$< 1.6 \times 10^{-5}$	CL=95%	1319
$f_0 \rightarrow K^+ K^-$			$\bar{\Lambda}_c^-(2593)^- / \bar{\Lambda}_c^-(2625)^- p$	$< 1.1 \times 10^{-4}$	CL=90%	—
$a_1(1260)^\mp \pi^\pm$	$[hh] (2.6 \pm 0.5) \times 10^{-5}$	S=1.9 2494	$\Xi_c^- \Lambda_c^+, \Xi_c^- \rightarrow \Xi^+ \pi^- \pi^-$	$(1.7 \pm 1.8) \times 10^{-5}$	S=2.2	1147
$a_2(1320)^\mp \pi^\pm$	$[hh] < 6.3 \times 10^{-6}$	CL=90%	$\Lambda_c^+ \Lambda_c^- K^0$	$(4.3 \pm 2.2) \times 10^{-4}$		—
$\pi^+ \pi^- \pi^0 \pi^0$	$< 3.1 \times 10^{-3}$	CL=90%	<b>Lepton Family number (LF) or Lepton number (L) or Baryon number (B) violating modes, or/and <math>\Delta B = 1</math> weak neutral current (B1) modes</b>			
$\rho^+ \rho^-$	$(2.77 \pm 0.19) \times 10^{-5}$	2523	$\gamma \gamma$	$B1 < 3.2 \times 10^{-7}$	CL=90%	2640
$a_1(1260)^0 \pi^0$	$< 1.1 \times 10^{-3}$	CL=90%	$e^+ e^-$	$B1 < 8.3 \times 10^{-8}$	CL=90%	2640
$\omega \pi^0$	$< 5 \times 10^{-7}$	CL=90%	$e^+ e^- \gamma$	$B1 < 1.2 \times 10^{-7}$	CL=90%	2640
$\pi^+ \pi^+ \pi^- \pi^- \pi^0$	$< 9.0 \times 10^{-3}$	CL=90%	$\mu^+ \mu^-$	$B1 (3.9 \pm 1.6) \times 10^{-10}$		2638
$a_1(1260)^+ \rho^-$	$< 6.1 \times 10^{-5}$	CL=90%	$\mu^+ \mu^- \gamma$	$B1 < 1.6 \times 10^{-7}$	CL=90%	2638
$a_1(1260)^0 \rho^0$	$< 2.4 \times 10^{-3}$	CL=90%				
$b_1^\mp \pi^\pm, b_1^\mp \rightarrow \omega \pi^\mp$	$(1.09 \pm 0.15) \times 10^{-5}$	—				
$b_1^0 \pi^0, b_1^0 \rightarrow \omega \pi^0$	$< 1.9 \times 10^{-6}$	CL=90%				
$b_1^- \rho^+, b_1^- \rightarrow \omega \pi^-$	$< 1.4 \times 10^{-6}$	CL=90%				
$b_1^0 \rho^0, b_1^0 \rightarrow \omega \pi^0$	$< 3.4 \times 10^{-6}$	CL=90%				

# Meson Summary Table

$\mu^+ \mu^- \mu^+ \mu^-$	$B1$	$< 5.3$	$\times 10^{-9}$	CL=90%	2629	
$SP, S \rightarrow \mu^+ \mu^-,$	$B1$	$[hh\alpha\alpha]$	$< 5.1$	$\times 10^{-9}$	CL=90%	—
$P \rightarrow \mu^+ \mu^-$						
$\tau^+ \tau^-$	$B1$	$< 4.1$	$\times 10^{-3}$	CL=90%	1952	
$\pi^0 \ell^+ \ell^-$	$B1$	$< 5.3$	$\times 10^{-8}$	CL=90%	2638	
$\pi^0 e^+ e^-$	$B1$	$< 8.4$	$\times 10^{-8}$	CL=90%	2638	
$\pi^0 \mu^+ \mu^-$	$B1$	$< 6.9$	$\times 10^{-8}$	CL=90%	2634	
$\eta \ell^+ \ell^-$	$B1$	$< 6.4$	$\times 10^{-8}$	CL=90%	2611	
$\eta e^+ e^-$	$B1$	$< 1.08$	$\times 10^{-7}$	CL=90%	2611	
$\eta \mu^+ \mu^-$	$B1$	$< 1.12$	$\times 10^{-7}$	CL=90%	2607	
$\pi^0 \nu \overline{\nu}$	$B1$	$< 6.9$	$\times 10^{-5}$	CL=90%	2638	
$K^0 \ell^+ \ell^-$	$B1$	$[ttt]$	$(3.1 \pm \frac{0.8}{0.7}) \times 10^{-7}$		2616	
$K^0 e^+ e^-$	$B1$	$(1.6 \pm \frac{1.0}{0.8}) \times 10^{-7}$			2616	
$K^0 \mu^+ \mu^-$	$B1$	$(3.39 \pm 0.34) \times 10^{-7}$			2612	
$K^0 \nu \overline{\nu}$	$B1$	$< 4.9$	$\times 10^{-5}$	CL=90%	2616	
$\rho^0 \nu \overline{\nu}$	$B1$	$< 2.08$	$\times 10^{-4}$	CL=90%	2583	
$K^*(892)^0 \ell^+ \ell^-$	$B1$	$[ttt]$	$(9.9 \pm \frac{1.2}{1.1}) \times 10^{-7}$		2565	
$K^*(892)^0 e^+ e^-$	$B1$	$(1.03 \pm \frac{0.19}{0.17}) \times 10^{-6}$			2565	
$K^*(892)^0 \mu^+ \mu^-$	$B1$	$(1.02 \pm 0.09) \times 10^{-6}$			2560	
$\pi^+ \pi^- \mu^+ \mu^-$		$(2.1 \pm 0.5) \times 10^{-8}$			2626	
$K^*(892)^0 \nu \overline{\nu}$	$B1$	$< 5.5$	$\times 10^{-5}$	CL=90%	2565	
$\phi \nu \overline{\nu}$	$B1$	$< 1.27$	$\times 10^{-4}$	CL=90%	2541	
$e^\pm \mu^\mp$	$LF$	$[hh]$	$< 2.8$	$\times 10^{-9}$	CL=90%	2639
$\pi^0 e^\pm \mu^\mp$	$LF$	$< 1.4$	$\times 10^{-7}$	CL=90%	2637	
$K^0 e^\pm \mu^\mp$	$LF$	$< 2.7$	$\times 10^{-7}$	CL=90%	2615	
$K^*(892)^0 e^+ \mu^-$	$LF$	$< 5.3$	$\times 10^{-7}$	CL=90%	2563	
$K^*(892)^0 e^- \mu^+$	$LF$	$< 3.4$	$\times 10^{-7}$	CL=90%	2563	
$K^*(892)^0 e^\pm \mu^\mp$	$LF$	$< 5.8$	$\times 10^{-7}$	CL=90%	2563	
$e^\pm \tau^\mp$	$LF$	$[hh]$	$< 2.8$	$\times 10^{-5}$	CL=90%	2341
$\mu^\pm \tau^\mp$	$LF$	$[hh]$	$< 2.2$	$\times 10^{-5}$	CL=90%	2339
invisible	$B1$	$< 2.4$	$\times 10^{-5}$	CL=90%	—	
$\nu \overline{\nu} \gamma$	$B1$	$< 1.7$	$\times 10^{-5}$	CL=90%	2640	
$\Lambda_c^+ \mu^-$	$L, B$	$< 1.4$	$\times 10^{-6}$	CL=90%	2143	
$\Lambda_c^+ e^-$	$L, B$	$< 4$	$\times 10^{-6}$	CL=90%	2145	

## $B^\pm/B^0$ ADMIXTURE

### CP violation

$$\begin{aligned}
 A_{CP}(B \rightarrow K^*(892)\gamma) &= -0.003 \pm 0.017 \\
 A_{CP}(b \rightarrow s\gamma) &= 0.015 \pm 0.020 \\
 A_{CP}(b \rightarrow (s+d)\gamma) &= 0.010 \pm 0.031 \\
 A_{CP}(B \rightarrow X_s \ell^+ \ell^-) &= 0.04 \pm 0.11 \\
 A_{CP}(B \rightarrow X_s \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4) &= -0.06 \pm 0.22 \\
 A_{CP}(B \rightarrow X_s \ell^+ \ell^-) (10.1 < q^2 < 12.9 \text{ or } q^2 > 14.2 \text{ GeV}^2/c^4) &= 0.19 \pm 0.18 \\
 A_{CP}(B \rightarrow K^* e^+ e^-) &= -0.18 \pm 0.15 \\
 A_{CP}(B \rightarrow K^* \mu^+ \mu^-) &= -0.03 \pm 0.13 \\
 A_{CP}(B \rightarrow K^* \ell^+ \ell^-) &= -0.04 \pm 0.07 \\
 A_{CP}(B \rightarrow \eta \text{ anything}) &= -0.13 \pm 0.04 \\
 \Delta A_{CP}(X_s \gamma) = A_{CP}(B^\pm \rightarrow X_s \gamma) - A_{CP}(B^0 \rightarrow X_s \gamma) &= 0.05 \pm 0.04
 \end{aligned}$$

The branching fraction measurements are for an admixture of  $B$  mesons at the  $\Upsilon(4S)$ . The values quoted assume that  $B(\Upsilon(4S) \rightarrow B\bar{B}) = 100\%$ .

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm \text{ anything}$ , the treatment of multiple  $D$ 's in the final state must be defined. One possibility would be to count the number of events with one-or-more  $D$ 's and divide by the total number of  $B$ 's. Another possibility would be to count the total number of  $D$ 's and divide by the total number of  $B$ 's, which is the definition of average multiplicity. The two definitions are identical if only one  $D$  is allowed in the final state. Even though the "one-or-more" definition seems sensible, for practical reasons inclusive branching fractions are almost always measured using the multiplicity definition. For heavy final state particles, authors call their results inclusive branching fractions while for light particles some authors call their results multiplicities. In the  $B$  sections, we list all results as inclusive branching fractions, adopting a multiplicity definition. This means that inclusive branching fractions can exceed 100% and that inclusive partial widths can exceed total widths, just as inclusive cross sections can exceed total cross section.

$\bar{B}$  modes are charge conjugates of the modes below. Reactions indicate the weak decay vertex and do not include mixing.

## B DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level (MeV/c)	$p$
<b>Semileptonic and leptonic modes</b>			
$\ell^+ \nu_\ell \text{ anything}$	$[ttt, iiaa]$	$(10.86 \pm 0.16) \%$	—
$D^- \ell^+ \nu_\ell \text{ anything}$	$[ttt]$	$(2.8 \pm 0.9) \%$	—
$\bar{D}^0 \ell^+ \nu_\ell \text{ anything}$	$[ttt]$	$(7.3 \pm 1.5) \%$	—
$\bar{D} \ell^+ \nu_\ell$		$(2.42 \pm 0.12) \%$	2310
$D^{*-} \ell^+ \nu_\ell \text{ anything}$	$[jja\bar{a}]$	$(6.7 \pm 1.3) \times 10^{-3}$	—
$D^* \ell^+ \nu_\ell$	$[kka\bar{a}]$	$(4.95 \pm 0.11) \%$	2257
$\bar{D}^{*+} \ell^+ \nu_\ell$	$[ttt, iiaa]$	$(2.7 \pm 0.7) \%$	—
$\bar{D}_1(2420) \ell^+ \nu_\ell \text{ anything}$		$(3.8 \pm 1.3) \times 10^{-3}$	S=2.4
$D \pi \ell^+ \nu_\ell \text{ anything} + D^* \pi \ell^+ \nu_\ell \text{ anything}$		$(2.6 \pm 0.5) \%$	S=1.5
$D \pi \ell^+ \nu_\ell \text{ anything}$		$(1.5 \pm 0.6) \%$	—
$D^* \pi \ell^+ \nu_\ell \text{ anything}$		$(1.9 \pm 0.4) \%$	—
$\bar{D}_2^*(2460) \ell^+ \nu_\ell \text{ anything}$		$(4.4 \pm 1.6) \times 10^{-3}$	—
$D^{*-} \pi^+ \ell^+ \nu_\ell \text{ anything}$		$(1.00 \pm 0.34) \%$	—
$\bar{D} \pi^+ \pi^- \ell^+ \nu_\ell$		$(1.62 \pm 0.32) \times 10^{-3}$	2301
$\bar{D}^* \pi^+ \pi^- \ell^+ \nu_\ell$		$(9.4 \pm 3.2) \times 10^{-4}$	2247
$D_s^- \ell^+ \nu_\ell \text{ anything}$	$[ttt]$	$< 7 \times 10^{-3}$	CL=90%
$D_s^- \ell^+ \nu_\ell K^+ \text{ anything}$	$[ttt]$	$< 5 \times 10^{-3}$	CL=90%
$D_s^- \ell^+ \nu_\ell K^0 \text{ anything}$	$[ttt]$	$< 7 \times 10^{-3}$	CL=90%
$X_c \ell^+ \nu_\ell$		$(10.65 \pm 0.16) \%$	—
$X_{u\ell} \ell^+ \nu_\ell$		$(2.14 \pm 0.31) \times 10^{-3}$	—
$K^+ \ell^+ \nu_\ell \text{ anything}$	$[ttt]$	$(6.3 \pm 0.6) \%$	—
$K^- \ell^+ \nu_\ell \text{ anything}$	$[ttt]$	$(10 \pm 4) \times 10^{-3}$	—
$K^0/\bar{K}^0 \ell^+ \nu_\ell \text{ anything}$	$[ttt]$	$(4.6 \pm 0.5) \%$	—
$\bar{D} \tau^+ \nu_\tau$		$(9.8 \pm 1.3) \times 10^{-3}$	1911
$D^* \tau^+ \nu_\tau$		$(1.58 \pm 0.12) \%$	1837

### $D, D^*, \text{ or } D_s$ modes

$D^\pm \text{ anything}$		$(22.9 \pm 1.3) \%$	—
$D^0/\bar{D}^0 \text{ anything}$		$(61.8 \pm 2.9) \%$	S=1.3
$D^*(2010)^\pm \text{ anything}$		$(22.5 \pm 1.5) \%$	—
$D^*(2007)^0 \text{ anything}$		$(26.0 \pm 2.7) \%$	—
$D_s^\pm \text{ anything}$	$[hh]$	$(8.3 \pm 0.8) \%$	—
$D_s^{*\pm} \text{ anything}$		$(6.3 \pm 1.0) \%$	—
$D_s^\pm \bar{D}^*(*)$		$(3.4 \pm 0.6) \%$	—
$\bar{D} D_{s0}(2317)$	seen		1605
$\bar{D} D_{sJ}(2457)$	seen		—
$D(*) \bar{D}^*(*) K^0 + D^*(*) \bar{D}^*(*) K^\pm$	$[hh, nnaa]$	$(7.1 \pm 2.7)$	—
$b \rightarrow c \bar{c} s$		$(22 \pm 4) \%$	—
$D_s^*(*) \bar{D}^*(*)$	$[hh, nnaa]$	$(3.9 \pm 0.4) \%$	—
$D^* D^*(2010)^\pm$	$[hh]$	$< 5.9 \times 10^{-3}$	CL=90%
$D D^*(2010)^\pm + D^* D^\pm$	$[hh]$	$< 5.5 \times 10^{-3}$	CL=90%
$D D^\pm$	$[hh]$	$< 3.1 \times 10^{-3}$	CL=90%
$D_s^*(*) \bar{D}^*(*) X(n\pi^\pm)$	$[hh, nnaa]$	$(9 \pm 5)$	—
$D^*(2010) \gamma$		$< 1.1 \times 10^{-3}$	CL=90%
$D_s^+ \pi^-, D_s^{*+} \pi^-, D_s^+ \rho^-, D_s^{*+} \rho^-, D_s^+ \eta, D_s^{*+} \eta, D_s^+ \rho^0, D_s^{*+} \rho^0, D_s^+ \omega, D_s^{*+} \omega$	$[hh]$	$< 4 \times 10^{-4}$	CL=90%
$D_{s1}(2536)^+ \text{ anything}$		$< 9.5 \times 10^{-3}$	CL=90%

### Charmonium modes

$J/\psi(1S) \text{ anything}$		$(1.094 \pm 0.032) \%$	S=1.1
$J/\psi(1S) \text{ (direct) anything}$		$(7.8 \pm 0.4) \times 10^{-3}$	S=1.1
$\psi(2S) \text{ anything}$		$(3.07 \pm 0.21) \times 10^{-3}$	—
$\chi_{c1}(1P) \text{ anything}$		$(3.86 \pm 0.27) \times 10^{-3}$	—
$\chi_{c1}(1P) \text{ (direct) anything}$		$(3.24 \pm 0.25) \times 10^{-3}$	—
$\chi_{c2}(1P) \text{ anything}$		$(1.4 \pm 0.4) \times 10^{-3}$	S=1.9
$\chi_{c2}(1P) \text{ (direct) anything}$		$(1.65 \pm 0.31) \times 10^{-3}$	—
$\eta_c(1S) \text{ anything}$		$< 9 \times 10^{-3}$	CL=90%
$KX(3872), X \rightarrow D^0 \bar{D}^0 \pi^0$		$(1.2 \pm 0.4) \times 10^{-4}$	1141
$KX(3872), X \rightarrow D^{*0} D^0$		$(8.0 \pm 2.2) \times 10^{-5}$	1141
$KX(3940), X \rightarrow D^{*0} D^0$		$< 6.7 \times 10^{-5}$	CL=90%
$KX(3915), X \rightarrow \omega J/\psi$	$[ooaa]$	$(7.1 \pm 3.4) \times 10^{-5}$	1103

### $K \text{ or } K^*$ modes

$K^\pm \text{ anything}$	$[hh]$	$(78.9 \pm 2.5) \%$	—
$K^+ \text{ anything}$		$(66 \pm 5) \%$	—
$K^- \text{ anything}$		$(13 \pm 4) \%$	—
$K^0/\bar{K}^0 \text{ anything}$	$[hh]$	$(64 \pm 4) \%$	—
$K^*(892)^\pm \text{ anything}$		$(18 \pm 6) \%$	—

## Meson Summary Table

$K^*(892)^0/\bar{K}^*(892)^0$ anything [hh]	( 14.6 ± 2.6 ) %	—
$K^*(892)\gamma$	( 4.2 ± 0.6 ) × 10 <sup>-5</sup>	2565
$\eta K\gamma$	( 8.5 ± 1.8 ) × 10 <sup>-6</sup>	2588
$K_1(1400)\gamma$	< 1.27 × 10 <sup>-4</sup>	CL=90% 2454
$K_2^*(1430)\gamma$	( 1.7 ± 0.6 ) × 10 <sup>-5</sup>	2447
$K_2(1770)\gamma$	< 1.2 × 10 <sup>-3</sup>	CL=90% 2342
$K_3^*(1780)\gamma$	< 3.7 × 10 <sup>-5</sup>	CL=90% 2341
$K_4^*(2045)\gamma$	< 1.0 × 10 <sup>-3</sup>	CL=90% 2244
$K\eta'(958)$	( 8.3 ± 1.1 ) × 10 <sup>-5</sup>	2528
$K^*(892)\eta'(958)$	( 4.1 ± 1.1 ) × 10 <sup>-6</sup>	2472
$K\eta$	< 5.2 × 10 <sup>-6</sup>	CL=90% 2588
$K^*(892)\eta$	( 1.8 ± 0.5 ) × 10 <sup>-5</sup>	2534
$K\phi\phi$	( 2.3 ± 0.9 ) × 10 <sup>-6</sup>	2306
$\bar{b} \rightarrow \bar{s}\gamma$	( 3.49 ± 0.19 ) × 10 <sup>-4</sup>	—
$\bar{b} \rightarrow \bar{d}\gamma$	( 9.2 ± 3.0 ) × 10 <sup>-6</sup>	—
$\bar{b} \rightarrow \bar{s}\text{gluon}$	< 6.8 %	CL=90% —
$\eta$ anything	( 2.6 ± 0.5 ) × 10 <sup>-4</sup>	—
$\eta'$ anything	( 4.2 ± 0.9 ) × 10 <sup>-4</sup>	—
$K^+$ gluon (charmless)	< 1.87 × 10 <sup>-4</sup>	CL=90% —
$K^0$ gluon (charmless)	( 1.9 ± 0.7 ) × 10 <sup>-4</sup>	—

## Light unflavored meson modes

$\rho\gamma$	( 1.39 ± 0.25 ) × 10 <sup>-6</sup>	S=1.2 2583
$\rho/\omega\gamma$	( 1.30 ± 0.23 ) × 10 <sup>-6</sup>	S=1.2 —
$\pi^\pm$ anything [hh,ppaa]	( 358 ± 7 ) %	—
$\pi^0$ anything	( 235 ± 11 ) %	—
$\eta$ anything	( 17.6 ± 1.6 ) %	—
$\rho^0$ anything	( 21 ± 5 ) %	—
$\omega$ anything	< 81 %	CL=90% —
$\phi$ anything	( 3.43 ± 0.12 ) %	—
$\phi K^*(892)$	< 2.2 × 10 <sup>-5</sup>	CL=90% 2460
$\pi^+$ gluon (charmless)	( 3.7 ± 0.8 ) × 10 <sup>-4</sup>	—

## Baryon modes

$\Lambda_c^+/\bar{\Lambda}_c^-$ anything	( 3.5 ± 0.4 ) %	—
$\Lambda_c^+$ anything	< 1.3 %	CL=90% —
$\bar{\Lambda}_c^-$ anything	< 7 %	CL=90% —
$\bar{\Lambda}_c^- \ell^+$ anything	< 9 × 10 <sup>-4</sup>	CL=90% —
$\bar{\Lambda}_c^- e^+$ anything	< 1.8 × 10 <sup>-3</sup>	CL=90% —
$\bar{\Lambda}_c^- \mu^+$ anything	< 1.4 × 10 <sup>-3</sup>	CL=90% —
$\bar{\Lambda}_c^- p$ anything	( 2.02 ± 0.33 ) %	—
$\bar{\Lambda}_c^- p e^+ \nu_e$	< 8 × 10 <sup>-4</sup>	CL=90% 2021
$\bar{\Sigma}_c^{--}$ anything	( 3.3 ± 1.7 ) × 10 <sup>-3</sup>	—
$\bar{\Sigma}_c^-$ anything	< 8 × 10 <sup>-3</sup>	CL=90% —
$\bar{\Sigma}_c^0$ anything	( 3.6 ± 1.7 ) × 10 <sup>-3</sup>	—
$\bar{\Sigma}_c^0 N(N=p \text{ or } n)$	< 1.2 × 10 <sup>-3</sup>	CL=90% 1938
$\Xi_c^0$ anything, $\Xi_c^0 \rightarrow \Xi^- \pi^+$	( 1.93 ± 0.30 ) × 10 <sup>-4</sup>	S=1.1 —
$\Xi_c^+, \Xi_c^+ \rightarrow \Xi^- \pi^+ \pi^+$	( 4.5 ± 1.3 ) × 10 <sup>-4</sup>	—
$p/\bar{p}$ anything [hh]	( 8.0 ± 0.4 ) %	—
$p/\bar{p}$ (direct) anything [hh]	( 5.5 ± 0.5 ) %	—
$\bar{p} e^+ \nu_e$ anything	< 5.9 × 10 <sup>-4</sup>	CL=90% —
$\Lambda/\bar{\Lambda}$ anything [hh]	( 4.0 ± 0.5 ) %	—
$\Lambda$ anything	seen	—
$\bar{\Lambda}$ anything	seen	—
$\Xi^-/\bar{\Xi}^+$ anything [hh]	( 2.7 ± 0.6 ) × 10 <sup>-3</sup>	—
baryons anything	( 6.8 ± 0.6 ) %	—
$p\bar{p}$ anything	( 2.47 ± 0.23 ) %	—
$\Lambda\bar{p}/\bar{\Lambda}p$ anything [hh]	( 2.5 ± 0.4 ) %	—
$\Lambda\bar{\Lambda}$ anything	< 5 × 10 <sup>-3</sup>	CL=90% —

Lepton Family number (LF) violating modes or  $\Delta B = 1$  weak neutral current ( $B1$ ) modes

$s e^+ e^-$	$B1$	( 6.7 ± 1.7 ) × 10 <sup>-6</sup>	S=2.0 —
$s \mu^+ \mu^-$	$B1$	( 4.3 ± 1.0 ) × 10 <sup>-6</sup>	—
$s \ell^+ \ell^-$	$B1$ [ttt]	( 5.8 ± 1.3 ) × 10 <sup>-6</sup>	S=1.8 —
$\pi \ell^+ \ell^-$	$B1$	< 5.9 × 10 <sup>-8</sup>	CL=90% 2638
$\pi e^+ e^-$	$B1$	< 1.10 × 10 <sup>-7</sup>	CL=90% 2638
$\pi \mu^+ \mu^-$	$B1$	< 5.0 × 10 <sup>-8</sup>	CL=90% 2634
$K e^+ e^-$	$B1$	( 4.4 ± 0.6 ) × 10 <sup>-7</sup>	2617
$K^*(892) e^+ e^-$	$B1$	( 1.19 ± 0.20 ) × 10 <sup>-6</sup>	S=1.2 2565
$K \mu^+ \mu^-$	$B1$	( 4.4 ± 0.4 ) × 10 <sup>-7</sup>	2612
$K^*(892) \mu^+ \mu^-$	$B1$	( 1.06 ± 0.09 ) × 10 <sup>-6</sup>	2560
$K \ell^+ \ell^-$	$B1$	( 4.8 ± 0.4 ) × 10 <sup>-7</sup>	2617

$K^*(892) \ell^+ \ell^-$	$B1$	( 1.05 ± 0.10 ) × 10 <sup>-6</sup>	2565
$K \nu \bar{\nu}$	$B1$	< 1.7 × 10 <sup>-5</sup>	CL=90% 2617
$K^* \nu \bar{\nu}$	$B1$	< 7.6 × 10 <sup>-5</sup>	CL=90% —
$s e^\pm \mu^\mp$	$LF$ [hh]	< 2.2 × 10 <sup>-5</sup>	CL=90% —
$\pi e^\pm \mu^\mp$	$LF$	< 9.2 × 10 <sup>-8</sup>	CL=90% 2637
$\rho e^\pm \mu^\mp$	$LF$	< 3.2 × 10 <sup>-6</sup>	CL=90% 2582
$K e^\pm \mu^\mp$	$LF$	< 3.8 × 10 <sup>-8</sup>	CL=90% 2616
$K^*(892) e^\pm \mu^\mp$	$LF$	< 5.1 × 10 <sup>-7</sup>	CL=90% 2563

 $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE

These measurements are for an admixture of bottom particles at high energy (LHC, LEP, Tevatron,  $Sp\bar{p}S$ ).

$$\text{Mean life } \tau = (1.566 \pm 0.003) \times 10^{-12} \text{ s}$$

$$\text{Mean life } \tau = (1.72 \pm 0.10) \times 10^{-12} \text{ s} \quad \text{Charged } b\text{-hadron admixture}$$

$$\text{Mean life } \tau = (1.58 \pm 0.14) \times 10^{-12} \text{ s} \quad \text{Neutral } b\text{-hadron admixture}$$

$$\tau^{\text{charged } b\text{-hadron}}/\tau^{\text{neutral } b\text{-hadron}} = 1.09 \pm 0.13$$

$$|\Delta\tau_b|/\tau_{b,\bar{b}} = -0.001 \pm 0.014$$

$$\text{Re}(\epsilon_b) / (1 + |\epsilon_b|^2) = (1.2 \pm 0.4) \times 10^{-3}$$

The branching fraction measurements are for an admixture of  $B$  mesons and baryons at energies above the  $T(4S)$ . Only the highest energy results (LHC, LEP, Tevatron,  $Sp\bar{p}S$ ) are used in the branching fraction averages. In the following, we assume that the production fractions are the same at the LHC, LEP, and at the Tevatron.

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm$  anything, the values usually are multiplicities, not branching fractions. They can be greater than one.

The modes below are listed for a  $\bar{b}$  initial state.  $b$  modes are their charge conjugates. Reactions indicate the weak decay vertex and do not include mixing.

 $\bar{B}$  DECAY MODES

Scale factor/  
Confidence level  $p$   
(MeV/c)

## PRODUCTION FRACTIONS

The production fractions for weakly decaying  $b$ -hadrons at high energy have been calculated from the best values of mean lives, mixing parameters, and branching fractions in this edition by the Heavy Flavor Averaging Group (HFAG) as described in the note “ $B^0$ - $\bar{B}^0$  Mixing” in the  $B^0$  Particle Listings. The production fractions in  $b$ -hadronic  $Z$  decay or  $p\bar{p}$  collisions at the Tevatron are also listed at the end of the section. Values assume

$$B(\bar{b} \rightarrow B^+) = B(\bar{b} \rightarrow B^0)$$

$$B(\bar{b} \rightarrow B^+) + B(\bar{b} \rightarrow B^0) + B(\bar{b} \rightarrow B_s^0) + B(b \rightarrow b\text{-baryon}) = 100\%.$$

The correlation coefficients between production fractions are also reported:

$$\text{cor}(B_s^0, b\text{-baryon}) = -0.240$$

$$\text{cor}(B_s^0, B^\pm=B^0) = -0.161$$

$$\text{cor}(b\text{-baryon}, B^\pm=B^0) = -0.920.$$

The notation for production fractions varies in the literature ( $f_d$ ,  $d_{B^0}$ ,  $f(b \rightarrow \bar{B}^0)$ ,  $\text{Br}(b \rightarrow \bar{B}^0)$ ). We use our own branching fraction notation here,  $B(\bar{b} \rightarrow B^0)$ .

Note these production fractions are  $b$ -hadronization fractions, not the conventional branching fractions of  $b$ -quark to a  $B$ -hadron, which may have considerable dependence on the initial and final state kinematic and production environment.

$B^+$	( 40.4 ± 0.6 ) %	—
$B^0$	( 40.4 ± 0.6 ) %	—
$B_s^0$	( 10.3 ± 0.5 ) %	—
$b$ -baryon	( 8.9 ± 1.3 ) %	—

## DECAY MODES

## Semileptonic and leptonic modes

$\nu$ anything	( 23.1 ± 1.5 ) %	—
$\ell^+ \nu_\ell$ anything [ttt]	( 10.69 ± 0.22 ) %	—
$e^+ \nu_e$ anything	( 10.86 ± 0.35 ) %	—
$\mu^+ \nu_\mu$ anything	( 10.95 ± 0.29 ) %	—
$D^- \ell^+ \nu_\ell$ anything [ttt]	( 2.2 ± 0.4 ) %	S=1.9 —
$D^- \pi^+ \ell^+ \nu_\ell$ anything	( 4.9 ± 1.9 ) × 10 <sup>-3</sup>	—
$D^- \pi^- \ell^+ \nu_\ell$ anything	( 2.6 ± 1.6 ) × 10 <sup>-3</sup>	—
$\bar{D}^0 \ell^+ \nu_\ell$ anything [ttt]	( 6.81 ± 0.34 ) %	—
$\bar{D}^0 \pi^- \ell^+ \nu_\ell$ anything	( 1.07 ± 0.27 ) %	—
$\bar{D}^0 \pi^+ \ell^+ \nu_\ell$ anything	( 2.3 ± 1.6 ) × 10 <sup>-3</sup>	—
$D^{*-} \ell^+ \nu_\ell$ anything [ttt]	( 2.75 ± 0.19 ) %	—

# Meson Summary Table

$D^{*-}\pi^-\ell^+\nu_\ell$ anything	( 6 ± 7 ) × 10 <sup>-4</sup>	-
$D^{*-}(pi+lepton)^+\nu_\ell$ anything	( 4.8 ± 1.0 ) × 10 <sup>-3</sup>	-
$\overline{D}_j^0\ell^+\nu_\ell$ anything × [ttt,qqaa]	( 2.6 ± 0.9 ) × 10 <sup>-3</sup>	-
$B(\overline{D}_j^0 \rightarrow D^{*+}\pi^-)$		-
$D_j^-\ell^+\nu_\ell$ anything × [ttt,qqaa]	( 7.0 ± 2.3 ) × 10 <sup>-3</sup>	-
$B(D_j^- \rightarrow D^0\pi^-)$		-
$\overline{D}_2^*(2460)^0\ell^+\nu_\ell$ anything × $B(\overline{D}_2^*(2460)^0 \rightarrow D^{*-}\pi^+)$	< 1.4 × 10 <sup>-3</sup> CL=90%	-
$D_2^*(2460)^-\ell^+\nu_\ell$ anything × $B(D_2^*(2460)^- \rightarrow D^0\pi^-)$	( 4.2 ± 1.5 ) × 10 <sup>-3</sup>	-
$\overline{D}_2^*(2460)^0\ell^+\nu_\ell$ anything × $B(\overline{D}_2^*(2460)^0 \rightarrow D^-\pi^+)$	( 1.6 ± 0.8 ) × 10 <sup>-3</sup>	-
charmless $\ell\overline{D}_\ell$	[ttt] ( 1.7 ± 0.5 ) × 10 <sup>-3</sup>	-
$\tau^+\nu_\tau$ anything	( 2.41 ± 0.23 ) %	-
$D^{*-}\tau\nu_\tau$ anything	( 9 ± 4 ) × 10 <sup>-3</sup>	-
$\overline{c} \rightarrow \ell^-\overline{\nu}_\ell$ anything	[ttt] ( 8.02 ± 0.19 ) %	-
$c \rightarrow \ell^+\nu_\ell$ anything	( 1.6 ± 0.4 ) %	-

## Charmed meson and baryon modes

$\overline{D}^0$ anything	( 59.0 ± 2.9 ) %	-
$D^0 D_s^\pm$ anything	[hh] ( 9.1 ± 4.0 ) %	-
$D^\mp D_s^\pm$ anything	[hh] ( 4.0 ± 2.3 ) %	-
$\overline{D}^0 D^0$ anything	[hh] ( 5.1 ± 2.0 ) %	-
$D^0 D^\pm$ anything	[hh] ( 2.7 ± 1.8 ) %	-
$D^\pm D^\mp$ anything	[hh] < 9 × 10 <sup>-3</sup> CL=90%	-
$D^-$ anything	( 22.5 ± 1.7 ) %	-
$D^*(2010)^+$ anything	( 17.3 ± 2.0 ) %	-
$D_1(2420)^0$ anything	( 5.0 ± 1.5 ) %	-
$D^*(2010)^\mp D_s^\pm$ anything	[hh] ( 3.3 ± 1.6 ) %	-
$D^0 D^*(2010)^\pm$ anything	[hh] ( 3.0 ± 1.1 ) %	-
$D^*(2010)^\pm D^\mp$ anything	[hh] ( 2.5 ± 1.2 ) %	-
$D^*(2010)^\pm D^*(2010)^\mp$ anything	[hh] ( 1.2 ± 0.4 ) %	-
$\overline{D} D$ anything	( 10 ± 11 ) %	-
$D_2^*(2460)^0$ anything	( 4.7 ± 2.7 ) %	-
$D_s^-$ anything	( 14.7 ± 2.1 ) %	-
$D_s^+$ anything	( 10.1 ± 3.1 ) %	-
$\Lambda_c^+$ anything	( 7.6 ± 1.1 ) %	-
$\overline{c}/c$ anything	[ppaa] (116.2 ± 3.2 ) %	-

## Charmonium modes

$J/\psi(1S)$ anything	( 1.16 ± 0.10 ) %	-
$\psi(2S)$ anything	( 2.83 ± 0.29 ) × 10 <sup>-3</sup>	-
$\chi_{c1}(1P)$ anything	( 1.4 ± 0.4 ) %	-

## K or K\* modes

$\overline{3}\gamma$	( 3.1 ± 1.1 ) × 10 <sup>-4</sup>	-
$\overline{3}\overline{D}\nu$	$B1$ < 6.4 × 10 <sup>-4</sup> CL=90%	-
$K^\pm$ anything	( 74 ± 6 ) %	-
$K_S^0$ anything	( 29.0 ± 2.9 ) %	-

## Pion modes

$\pi^\pm$ anything	(397 ± 21 ) %	-
$\pi^0$ anything	[ppaa] (278 ± 60 ) %	-
$\phi$ anything	( 2.82 ± 0.23 ) %	-

## Baryon modes

$p/\overline{p}$ anything	( 13.1 ± 1.1 ) %	-
$\Lambda/\overline{\Lambda}$ anything	( 5.9 ± 0.6 ) %	-
$b$ -baryon anything	( 10.2 ± 2.8 ) %	-

## Other modes

charged anything	[ppaa] (497 ± 7 ) %	-
hadron <sup>+</sup> hadron <sup>-</sup>	( 1.7 ± 1.0 ) × 10 <sup>-5</sup>	-
charmless	( 7 ± 21 ) × 10 <sup>-3</sup>	-

## $\Delta B = 1$ weak neutral current ( $B1$ ) modes

$\mu^+\mu^-$ anything	$B1$ < 3.2 × 10 <sup>-4</sup> CL=90%	-
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### $B^*$

$$I(J^P) = \frac{1}{2}(1^-)$$

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

$$\text{Mass } m_{B^*} = 5324.65 \pm 0.25 \text{ MeV}$$

$$m_{B^*} - m_B = 45.18 \pm 0.23 \text{ MeV}$$

$$m_{B^{*+}} - m_{B^+} = 45.34 \pm 0.23 \text{ MeV}$$

## $B^*$ DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
--	--------------------------------	----------------

$B\gamma$	dominant	45
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### $B_1(5721)^+$

$$I(J^P) = \frac{1}{2}(1^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 5725.9 \pm 2.5 \text{ MeV}$$

$$m_{B_1^+} - m_{B^{*0}} = 401.2 \pm 2.4 \text{ MeV}$$

$$\text{Full width } \Gamma = 31 \pm 6 \text{ MeV } (S = 1.1)$$

## $B_1(5721)^+$ DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
--	--------------------------------	----------------

$B^{*0}\pi^+$	seen	363
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### $B_1(5721)^0$

$$I(J^P) = \frac{1}{2}(1^+)$$

$I, J, P$  need confirmation.

$$B_1(5721)^0 \text{ MASS} = 5726.0 \pm 1.3 \text{ MeV } (S = 1.2)$$

$$m_{B_1^0} - m_{B^+} = 446.7 \pm 1.3 \text{ MeV } (S = 1.2)$$

$$m_{B_1^0} - m_{B^{*+}} = 401.4 \pm 1.2 \text{ MeV } (S = 1.2)$$

$$\text{Full width } \Gamma = 27.5 \pm 3.4 \text{ MeV } (S = 1.1)$$

## $B_1(5721)^0$ DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
--	--------------------------------	----------------

$B^{*+}\pi^-$	dominant	363
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### $B_2^*(5747)^+$

$$I(J^P) = \frac{1}{2}(2^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 5737.2 \pm 0.7 \text{ MeV}$$

$$m_{B_2^{*+}} - m_{B^0} = 457.5 \pm 0.7 \text{ MeV}$$

$$\text{Full width } \Gamma = 20 \pm 5 \text{ MeV } (S = 2.2)$$

## $B_2^*(5747)^+$ DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
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$B^0\pi^+$	seen	418
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$B^{*0}\pi^+$	seen	374
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### $B_2^*(5747)^0$

$$I(J^P) = \frac{1}{2}(2^+)$$

$I, J, P$  need confirmation.

$$B_2^*(5747)^0 \text{ MASS} = 5739.5 \pm 0.7 \text{ MeV } (S = 1.4)$$

$$m_{B_2^{*0}} - m_{B_1^0} = 13.5 \pm 1.4 \text{ MeV } (S = 1.3)$$

$$m_{B_2^{*0}} - m_{B^+} = 460.2 \pm 0.6 \text{ MeV } (S = 1.4)$$

$$\text{Full width } \Gamma = 24.2 \pm 1.7 \text{ MeV}$$

## $B_2^*(5747)^0$ DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
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$B^+\pi^-$	dominant	421
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$B^{*+}\pi^-$	dominant	376
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### $B_J(5970)^+$

$$I(J^P) = \frac{1}{2}(?)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 5964 \pm 5 \text{ MeV}$$

$$m_{B_J(5970)^+} - m_{B^0} = 685 \pm 5 \text{ MeV}$$

$$m_{B_J(5970)^+} - m_{B^{*0}}$$

$$\text{Full width } \Gamma = 62 \pm 20 \text{ MeV}$$

# Meson Summary Table

$B_J(5970)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$B^0 \pi^+$	possibly seen	632
$B^{*0} \pi^+$	seen	591

<b><math>B_J(5970)^0</math></b>	$I(J^P) = \frac{1}{2}(?)^?$ $I, J, P$ need confirmation.
Mass $m = 5971 \pm 5$ MeV	
$m_{B_J(5970)^0} - m_{B^+} = 691 \pm 5$ MeV	
$m_{B_J(5970)^0} - m_{B^{*+}}$	
Full width $\Gamma = 81 \pm 12$ MeV	

$B_J(5970)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$B^+ \pi^-$	possibly seen	638
$B^{*+} \pi^-$	seen	597

## BOTTOM, STRANGE MESONS ( $B = \pm 1, S = \mp 1$ )

$$B_s^0 = s\bar{b}, \bar{B}_s^0 = \bar{s}b, \text{ similarly for } B_s^{*\prime}s$$

<b><math>B_s^0</math></b>	$I(J^P) = 0(0^-)$
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$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

$$\text{Mass } m_{B_s^0} = 5366.82 \pm 0.22 \text{ MeV}$$

$$m_{B_s^0} - m_B = 87.35 \pm 0.20 \text{ MeV}$$

$$\text{Mean life } \tau = (1.510 \pm 0.005) \times 10^{-12} \text{ s}$$

$$c\tau = 452.7 \mu\text{m}$$

$$\Delta\Gamma_{B_s^0} = \Gamma_{B_{sL}^0} - \Gamma_{B_{sH}^0} = (0.082 \pm 0.007) \times 10^{12} \text{ s}^{-1}$$

### $B_s^0$ - $\bar{B}_s^0$ mixing parameters

$$\Delta m_{B_s^0} = m_{B_{sH}^0} - m_{B_{sL}^0} = (17.757 \pm 0.021) \times 10^{12} \hbar \text{ s}^{-1}$$

$$= (1.1688 \pm 0.0014) \times 10^{-8} \text{ MeV}$$

$$x_s = \Delta m_{B_s^0}/\Gamma_{B_s^0} = 26.81 \pm 0.10$$

$$\chi_s = 0.499308 \pm 0.000005$$

### CP violation parameters in $B_s^0$

$$\text{Re}(\epsilon_{B_s^0}) / (1 + |\epsilon_{B_s^0}|^2) = (-1.9 \pm 1.0) \times 10^{-3}$$

$$C_{KK}(B_s^0 \rightarrow K^+ K^-) = 0.14 \pm 0.11$$

$$S_{KK}(B_s^0 \rightarrow K^+ K^-) = 0.30 \pm 0.13$$

$$\gamma = (65 \pm 7)^\circ$$

$$\delta_B(B_s^0 \rightarrow D_s^\pm K^\mp) = (3 \pm 20)^\circ$$

$$r_B(B_s^0 \rightarrow D_s^\mp K^\pm) = 0.53 \pm 0.17$$

$$\text{CP Violation phase } \beta_s = (0.6 \pm 1.9) \times 10^{-2} \text{ rad}$$

$$|\lambda| (B_s^0 \rightarrow J/\psi(1S)\phi) = 0.964 \pm 0.020$$

$$|\lambda| = 1.02 \pm 0.07$$

$$A, \text{ CP violation parameter} = 0.5^{+0.8}_{-0.7}$$

$$C, \text{ CP violation parameter} = -0.3 \pm 0.4$$

$$S, \text{ CP violation parameter} = -0.1 \pm 0.4$$

$$A_{CP}^L(B_s \rightarrow J/\psi \bar{K}^*(892)^0) = -0.05 \pm 0.06$$

$$A_{CP}^|| (B_s \rightarrow J/\psi \bar{K}^*(892)^0) = 0.17 \pm 0.15$$

$$A_{CP}^{\perp}(B_s \rightarrow J/\psi \bar{K}^*(892)^0) = -0.05 \pm 0.10$$

$$A_{CP}(B_s \rightarrow \pi^+ K^-) = 0.263 \pm 0.035$$

$$A_{CP}(B_s^0 \rightarrow [K^+ K^-]_D \bar{K}^*(892)^0) = -0.04 \pm 0.07$$

$$A_{CP}(B_s^0 \rightarrow [\pi^+ K^-]_D K^*(892)^0) = -0.01 \pm 0.04$$

$$A_{CP}(B_s^0 \rightarrow [\pi^+ \pi^-]_D K^*(892)^0) = 0.06 \pm 0.13$$

$$\Delta a_{\perp} < 1.2 \times 10^{-12} \text{ GeV, CL} = 95\%$$

These branching fractions all scale with  $B(\bar{b} \rightarrow B_s^0)$ .

The branching fraction  $B(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything})$  is not a pure measurement since the measured product branching fraction  $B(\bar{b} \rightarrow B_s^0) \times B(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything})$  was used to determine  $B(\bar{b} \rightarrow B_s^0)$ , as described in the note on " $B^0$ - $\bar{B}^0$  Mixing"

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm \text{ anything}$ , the values usually are multiplicities, not branching fractions. They can be greater than one.

$B_s^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor / Confidence level	$\rho$ (MeV/c)
$D_s^-$ anything	(93 $\pm$ 25 ) %		—
$\ell \nu_\ell X$	( 9.6 $\pm$ 0.8 ) %		—
$e^+ \nu X^-$	( 9.1 $\pm$ 0.8 ) %		—
$\mu^+ \nu X^-$	(10.2 $\pm$ 1.0 ) %		—
$D_s^- \ell^+ \nu_\ell \text{ anything}$	[raa] ( 8.1 $\pm$ 1.3 ) %		—
$D_s^{*-} \ell^+ \nu_\ell \text{ anything}$	( 5.4 $\pm$ 1.1 ) %		—
$D_{s1}(2536)^- \mu^+ \nu_\mu$	( 2.6 $\pm$ 0.7 ) $\times 10^{-3}$		—
$D_{s1}^- \rightarrow D^{*-} K_S^0$			
$D_{s1}(2536)^- X \mu^+ \nu$	( 4.4 $\pm$ 1.3 ) $\times 10^{-3}$		—
$D_{s1}^- \rightarrow \bar{D}^0 K^+$			
$D_{s2}(2573)^- X \mu^+ \nu$	( 2.7 $\pm$ 1.0 ) $\times 10^{-3}$		—
$D_{s2}^- \rightarrow \bar{D}^0 K^+$			
$D_s^- \pi^+$	( 3.00 $\pm$ 0.23 ) $\times 10^{-3}$		2320
$D_s^- \rho^+$	( 6.9 $\pm$ 1.4 ) $\times 10^{-3}$		2249
$D_s^- \pi^+ \pi^+ \pi^-$	( 6.1 $\pm$ 1.0 ) $\times 10^{-3}$		2301
$D_{s1}(2536)^- \pi^+$	( 2.5 $\pm$ 0.8 ) $\times 10^{-5}$		—
$D_{s1}^- \rightarrow D_s^- \pi^+ \pi^-$			
$D_s^\mp K^\pm$	( 2.27 $\pm$ 0.19 ) $\times 10^{-4}$		2293
$D_s^{*-} K^+ \pi^+ \pi^-$	( 3.2 $\pm$ 0.6 ) $\times 10^{-4}$		2249
$D_s^+ D_s^-$	( 4.4 $\pm$ 0.5 ) $\times 10^{-3}$		1824
$D_s^- D^+$	( 2.8 $\pm$ 0.5 ) $\times 10^{-4}$		1875
$D^+ D^-$	( 2.2 $\pm$ 0.6 ) $\times 10^{-4}$		1925
$D^0 \bar{D}^0$	( 1.9 $\pm$ 0.5 ) $\times 10^{-4}$		1930
$D^{*-} \pi^+$	( 2.0 $\pm$ 0.5 ) $\times 10^{-3}$		2265
$D_s^{*\mp} K^\pm$	( 1.33 $\pm$ 0.35 ) $\times 10^{-4}$		—
$D_s^{*-} \rho^+$	( 9.6 $\pm$ 2.1 ) $\times 10^{-3}$		2191
$D_s^{*+} D^- + D_s^{*-} D^+$	( 1.29 $\pm$ 0.22 ) %	S=1.1	1742
$D_s^{*+} D_s^{*-}$	( 1.86 $\pm$ 0.30 ) %		1655
$D^{(*)+} D^{(*)-}$	( 4.5 $\pm$ 1.4 ) %		—
$\bar{D}^0 K^- \pi^+$	( 1.03 $\pm$ 0.13 ) $\times 10^{-3}$		2312
$\bar{D}^0 \bar{K}^*(892)^0$	( 4.4 $\pm$ 0.6 ) $\times 10^{-4}$		2264
$\bar{D}^0 \bar{K}^*(1410)$	( 3.9 $\pm$ 3.5 ) $\times 10^{-4}$		2117
$\bar{D}^0 \bar{K}_0^*(1430)$	( 3.0 $\pm$ 0.7 ) $\times 10^{-4}$		2113
$\bar{D}^0 \bar{K}_2^*(1430)$	( 1.1 $\pm$ 0.4 ) $\times 10^{-4}$		2113
$\bar{D}^0 \bar{K}^*(1680)$	< 7.8 $\times 10^{-5}$	CL=90%	1998
$\bar{D}^0 \bar{K}_0^*(1950)$	< 1.1 $\times 10^{-4}$	CL=90%	1890
$\bar{D}^0 \bar{K}_2^*(1780)$	< 2.6 $\times 10^{-5}$	CL=90%	1971
$\bar{D}^0 \bar{K}_4^*(2045)$	< 3.1 $\times 10^{-5}$	CL=90%	1837
$\bar{D}^0 K^- \pi^+$ (non-resonant)	( 2.1 $\pm$ 0.8 ) $\times 10^{-4}$		2312
$D_{s2}^-(2573) \pi^+$	( 2.6 $\pm$ 0.4 ) $\times 10^{-4}$		—
$D_{s2}^{*-} \rightarrow \bar{D}^0 K^-$			
$D_{s1}^*(2700) \pi^+$	( 1.6 $\pm$ 0.8 ) $\times 10^{-5}$		—
$D_{s1}^* \rightarrow \bar{D}^0 K^-$			
$D_{s1}^*(2860) \pi^+$	( 5 $\pm$ 4 ) $\times 10^{-5}$		—
$D_{s1}^* \rightarrow \bar{D}^0 K^-$			
$D_{s3}^*(2860) \pi^+$	( 2.2 $\pm$ 0.6 ) $\times 10^{-5}$		—
$D_{s3}^* \rightarrow \bar{D}^0 K^-$			
$\bar{D}^0 K^+ K^-$	( 4.4 $\pm$ 2.0 ) $\times 10^{-5}$		2243
$\bar{D}^0 f_0(980)$	< 3.1 $\times 10^{-6}$	CL=90%	2242
$\bar{D}^0 \phi$	( 3.0 $\pm$ 0.8 ) $\times 10^{-5}$		2235
$D^{*\mp} \pi^\pm$	< 6.1 $\times 10^{-6}$	CL=90%	—
$J/\psi(1S) \phi$	( 1.07 $\pm$ 0.08 ) $\times 10^{-3}$		1588
$J/\psi(1S) \pi^0$	< 1.2 $\times 10^{-3}$	CL=90%	1786
$J/\psi(1S) \eta$	( 3.9 $\pm$ 0.7 ) $\times 10^{-4}$	S=1.4	1733
$J/\psi(1S) K_S^0$	( 1.89 $\pm$ 0.12 ) $\times 10^{-5}$		1743
$J/\psi(1S) \bar{K}^*(892)^0$	( 4.1 $\pm$ 0.4 ) $\times 10^{-5}$		1637
$J/\psi(1S) \eta'$	( 3.3 $\pm$ 0.4 ) $\times 10^{-4}$		1612
$J/\psi(1S) \pi^+ \pi^-$	( 2.13 $\pm$ 0.18 ) $\times 10^{-4}$		1775
$J/\psi(1S) f_0(500), f_0 \rightarrow \pi^+ \pi^-$	< 1.7 $\times 10^{-6}$	CL=90%	—

# Meson Summary Table

$J/\psi(1S) \rho, \rho \rightarrow \pi^+ \pi^-$	$< 1.2$	$\times 10^{-6}$	CL=90%	—
$J/\psi(1S) f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$(1.34 \pm 0.15)$	$\times 10^{-4}$		—
$J/\psi(1S) f_0(980)_0, f_0 \rightarrow \pi^+ \pi^-$	$(5.1 \pm 0.9)$	$\times 10^{-5}$		—
$J/\psi(1S) f_2(1270)_0, f_2 \rightarrow \pi^+ \pi^-$	$(2.6 \pm 0.7)$	$\times 10^{-7}$		—
$J/\psi(1S) f_2(1270)_\parallel, f_2 \rightarrow \pi^+ \pi^-$	$(3.8 \pm 1.3)$	$\times 10^{-7}$		—
$J/\psi(1S) f_2(1270)_\perp, f_2 \rightarrow \pi^+ \pi^-$	$(4.6 \pm 2.7)$	$\times 10^{-7}$		—
$J/\psi(1S) f_0(1500), f_0 \rightarrow \pi^+ \pi^-$	$(7.3 \pm 1.6)$	$\times 10^{-6}$		—
$J/\psi(1S) f'_2(1525)_0, f'_2 \rightarrow \pi^+ \pi^-$	$(3.7 \pm 1.0)$	$\times 10^{-7}$		—
$J/\psi(1S) f'_2(1525)_\parallel, f'_2 \rightarrow \pi^+ \pi^-$	$(4.3 \pm 9.0)$	$\times 10^{-8}$		—
$J/\psi(1S) f'_2(1525)_\perp, f'_2 \rightarrow \pi^+ \pi^-$	$(1.9 \pm 1.4)$	$\times 10^{-7}$		—
$J/\psi(1S) f_0(1790), f_0 \rightarrow \pi^+ \pi^-$	$(1.7 \pm 4.0)$	$\times 10^{-6}$		—
$J/\psi(1S) \bar{K}^0 \pi^+ \pi^-$	$< 4.4$	$\times 10^{-5}$	CL=90%	1675
$J/\psi(1S) K^+ K^-$	$(7.9 \pm 0.7)$	$\times 10^{-4}$		1601
$J/\psi(1S) K^0 \bar{K}^- \pi^+ + \text{c.c.}$	$(9.3 \pm 1.3)$	$\times 10^{-4}$		1538
$J/\psi(1S) \bar{K}^0 K^+ K^-$	$< 1.2$	$\times 10^{-5}$	CL=90%	1333
$J/\psi(1S) f'_2(1525)$	$(2.6 \pm 0.6)$	$\times 10^{-4}$		1304
$J/\psi(1S) p \bar{p}$	$< 4.8$	$\times 10^{-6}$	CL=90%	982
$J/\psi(1S) \gamma$	$< 7.3$	$\times 10^{-6}$	CL=90%	1790
$J/\psi(1S) \pi^+ \pi^- \pi^+ \pi^-$	$(7.9 \pm 0.9)$	$\times 10^{-5}$		1731
$J/\psi(1S) f_1(1285)$	$(7.1 \pm 1.4)$	$\times 10^{-5}$		1460
$\psi(2S) \eta$	$(3.3 \pm 0.9)$	$\times 10^{-4}$		1338
$\psi(2S) \eta'$	$(1.29 \pm 0.35)$	$\times 10^{-4}$		1158
$\psi(2S) \pi^+ \pi^-$	$(7.2 \pm 1.2)$	$\times 10^{-5}$		1397
$\psi(2S) \phi$	$(5.4 \pm 0.5)$	$\times 10^{-4}$		1120
$\psi(2S) K^- \pi^+$	$(3.12 \pm 0.30)$	$\times 10^{-5}$		1310
$\psi(2S) \bar{K}^*(892)^0$	$(3.3 \pm 0.5)$	$\times 10^{-5}$		1196
$\chi_{c1} \phi$	$(2.03 \pm 0.29)$	$\times 10^{-4}$		1274
$\pi^+ \pi^-$	$(7.7 \pm 2.0)$	$\times 10^{-7}$	S=1.4	2680
$\pi^0 \pi^0$	$< 2.1$	$\times 10^{-4}$	CL=90%	2680
$\eta \pi^0$	$< 1.0$	$\times 10^{-3}$	CL=90%	2654
$\eta \eta$	$< 1.5$	$\times 10^{-3}$	CL=90%	2627
$\rho^0 \rho^0$	$< 3.20$	$\times 10^{-4}$	CL=90%	2569
$\eta' \eta'$	$(3.3 \pm 0.7)$	$\times 10^{-5}$		2507
$\phi \rho^0$	$< 6.17$	$\times 10^{-4}$	CL=90%	2526
$\phi \phi$	$(1.87 \pm 0.15)$	$\times 10^{-5}$		2482
$\pi^+ K^-$	$(5.6 \pm 0.6)$	$\times 10^{-6}$		2659
$K^+ K^-$	$(2.52 \pm 0.17)$	$\times 10^{-5}$		2638
$K^0 \bar{K}^0$	$< 6.6$	$\times 10^{-5}$	CL=90%	2637
$K^0 \pi^+ \pi^-$	$(1.5 \pm 0.4)$	$\times 10^{-5}$		2653
$K^0 K^\pm \pi^\mp$	$(7.7 \pm 1.0)$	$\times 10^{-5}$		2622
$K^*(892)^- \pi^+$	$(3.3 \pm 1.2)$	$\times 10^{-6}$		2607
$K^*(892)^\pm K^\mp$	$(1.25 \pm 0.26)$	$\times 10^{-5}$		2585
$K_S^0 \bar{K}^*(892)^0 + \text{c.c.}$	$(1.6 \pm 0.4)$	$\times 10^{-5}$		2585
$K_S^0 K^+ K^-$	$< 3.5$	$\times 10^{-6}$	CL=90%	2568
$\bar{K}^*(892)^0 \rho^0$	$< 7.67$	$\times 10^{-4}$	CL=90%	2550
$\bar{K}^*(892)^0 K^*(892)^0$	$(1.11 \pm 0.27)$	$\times 10^{-5}$		2531
$\phi K^*(892)^0$	$(1.14 \pm 0.30)$	$\times 10^{-6}$		2507
$p \bar{p}$	$(2.8 \pm 2.2)$	$\times 10^{-8}$		2514
$\Lambda_c^- \Lambda \pi^+$	$(3.6 \pm 1.6)$	$\times 10^{-4}$		—
$\Lambda_c^- \Lambda_c^+$	$< 8.0$	$\times 10^{-5}$	CL=95%	—
$\gamma \gamma$	$< 3.1$	$\times 10^{-6}$	CL=90%	2683
$\phi \gamma$	$(3.52 \pm 0.34)$	$\times 10^{-5}$		2587

## Lepton Family number (LF) violating modes or $\Delta B = 1$ weak neutral current (B1) modes

$\mu^+ \mu^-$	B1	$(2.9 \pm 0.7)$	$\times 10^{-9}$	2681
$e^+ e^-$	B1	$< 2.8$	$\times 10^{-7}$	CL=90% 2683
$\mu^+ \mu^- \mu^+ \mu^-$	B1	$< 1.2$	$\times 10^{-8}$	CL=90% 2673
$S P, S \rightarrow \mu^+ \mu^-, P \rightarrow \mu^+ \mu^-$	B1 [hhaa]	$< 1.2$	$\times 10^{-8}$	CL=90% —
$\phi(1020) \mu^+ \mu^-$	B1	$(8.2 \pm 1.2)$	$\times 10^{-7}$	2582

$\pi^+ \pi^- \mu^+ \mu^-$	B1	$(8.4 \pm 1.7)$	$\times 10^{-8}$	2670
$\phi \nu \bar{\nu}$	B1	$< 5.4$	$\times 10^{-3}$	CL=90% 2587
$e^\pm \mu^\mp$	LF [hh]	$< 1.1$	$\times 10^{-8}$	CL=90% 2682

## $B_s^*$

$$I(J^P) = 0(1^-)$$

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

$$\text{Mass } m = 5415.4^{+1.8}_{-1.5} \text{ MeV} \quad (S = 3.0)$$

$$m_{B_s^*} - m_{B_s} = 48.6^{+1.8}_{-1.6} \text{ MeV} \quad (S = 2.8)$$

$B_s^*$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$B_s \gamma$	dominant	—

## $B_{s1}(5830)^0$

$$I(J^P) = 0(1^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 5828.63 \pm 0.27 \text{ MeV}$$

$$m_{B_{s1}^0} - m_{B^{*+}} = 503.98 \pm 0.18 \text{ MeV}$$

$$\text{Full width } \Gamma = 0.5 \pm 0.4 \text{ MeV}$$

$B_{s1}(5830)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$B^{*+} K^-$	dominant	97

## $B_{s2}^*(5840)^0$

$$I(J^P) = 0(2^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 5839.84 \pm 0.18 \text{ MeV} \quad (S = 1.1)$$

$$m_{B_{s2}^0} - m_{B_{s1}^0}$$

$$m_{B_{s2}^0} - m_{B^{*+}} = 560.53 \pm 0.18 \text{ MeV} \quad (S = 1.1)$$

$$\text{Full width } \Gamma = 1.47 \pm 0.33 \text{ MeV}$$

$B_{s2}^*(5840)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$B^+ K^-$	dominant	253

# BOTTOM, CHARMED MESONS ( $B = C = \pm 1$ )

$$B_c^+ = c \bar{b}, B_c^- = \bar{c} b, \text{ similarly for } B_c^{* \text{'s}}$$

## $B_c^+$

$$I(J^P) = 0(0^-)$$

$I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

$$\text{Mass } m = 6275.1 \pm 1.0 \text{ MeV}$$

$$\text{Mean life } \tau = (0.507 \pm 0.009) \times 10^{-12} \text{ s}$$

$B_c^-$  modes are charge conjugates of the modes below.

$B_c^+$ DECAY MODES $\times B(\bar{b} \rightarrow B_c)$	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
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The following quantities are not pure branching ratios; rather the fraction  $\Gamma_i/\Gamma \times B(\bar{b} \rightarrow B_c)$ .

$J/\psi(1S) \ell^+ \nu_\ell \text{ anything}$	$(5.2 \pm 2.4)$	$\times 10^{-5}$	—
$J/\psi(1S) \pi^+$	seen		2371
$J/\psi(1S) K^+$	seen		2341
$J/\psi(1S) \pi^+ \pi^+ \pi^-$	seen		2350
$J/\psi(1S) a_1(1260)$	$< 1.2$	$\times 10^{-3}$	90% 2170
$J/\psi(1S) K^+ K^- \pi^+$	seen		2203
$J/\psi(1S) \pi^+ \pi^+ \pi^+ \pi^- \pi^-$	seen		2309
$\psi(2S) \pi^+$	seen		2052
$J/\psi(1S) D_s^+$	seen		1822
$J/\psi(1S) D_s^{*+}$	seen		1728
$J/\psi(1S) p \bar{p} \pi^+$	seen		1792
$D^*(2010)^+ \bar{D}^0$	$< 6.2$	$\times 10^{-3}$	90% 2467
$D^+ K^{*0}$	$< 0.20$	$\times 10^{-6}$	90% 2783

## Meson Summary Table

$D^+ \bar{K}^{*0}$	$< 0.16$	$\times 10^{-6}$	90%	2783
$D_s^+ K^{*0}$	$< 0.28$	$\times 10^{-6}$	90%	2751
$D_s^+ \bar{K}^{*0}$	$< 0.4$	$\times 10^{-6}$	90%	2751
$D_s^+ \phi$	$< 0.32$	$\times 10^{-6}$	90%	2727
$K^+ K^0$	$< 4.6$	$\times 10^{-7}$	90%	3098
$B_s^0 \pi^+ / B(\bar{D} \rightarrow B_s)$	$(2.37^{+0.37}_{-0.35}) \times 10^{-3}$			—

 **$c\bar{c}$  MESONS** **$\eta_c(1S)$** 

$$J^G(J^{PC}) = 0^+(0^{-+})$$

Mass  $m = 2983.4 \pm 0.5$  MeV ( $S = 1.2$ )

Full width  $\Gamma = 31.8 \pm 0.8$  MeV

$\eta_c(1S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
<b>Decays involving hadronic resonances</b>			
$\eta'(958) \pi \pi$	( 4.1 $\pm$ 1.7 ) %		1323
$\rho \rho$	( 1.8 $\pm$ 0.5 ) %		1274
$K^*(892)^0 K^- \pi^+ + c.c.$	( 2.0 $\pm$ 0.7 ) %		1277
$K^*(892) \bar{K}^*(892)$	( 7.0 $\pm$ 1.3 ) $\times 10^{-3}$		1196
$K^*(892)^0 \bar{K}^*(892)^0 \pi^+ \pi^-$	( 1.1 $\pm$ 0.5 ) %		1073
$\phi K^+ K^-$	( 2.9 $\pm$ 1.4 ) $\times 10^{-3}$		1104
$\phi \phi$	( 1.75 $\pm$ 0.20 ) $\times 10^{-3}$		1089
$\phi 2(\pi^+ \pi^-)$	$< 4$ $\times 10^{-3}$	90%	1251
$a_0(980) \pi$	$< 2$ %	90%	1327
$a_2(1320) \pi$	$< 2$ %	90%	1196
$K^*(892) \bar{K} + c.c.$	$< 1.28$ %	90%	1309
$f_2(1270) \eta$	$< 1.1$ %	90%	1145
$\omega \omega$	$< 3.1$ $\times 10^{-3}$	90%	1270
$\omega \phi$	$< 1.7$ $\times 10^{-3}$	90%	1185
$f_2(1270) f_2(1270)$	( 9.8 $\pm$ 2.5 ) $\times 10^{-3}$		774
$f_2(1270) f_2'(1525)$	( 9.7 $\pm$ 3.2 ) $\times 10^{-3}$		513
$f_0(980) \eta$	seen		1264
$f_0(1500) \eta$	seen		1026
$f_0(2200) \eta$	seen		496
$a_0(980) \pi$	seen		1327
$a_0(1320) \pi$	seen		—
$a_0(1450) \pi$	seen		1123
$a_0(1950) \pi$	seen		859
$a_2(1950) \pi$	not seen		—
$K_0^*(1430) \bar{K}$	seen		—
$K_2^*(1430) \bar{K}$	seen		—
$K_0^*(1950) \bar{K}$	seen		—
<b>Decays into stable hadrons</b>			
$K \bar{K} \pi$	( 7.3 $\pm$ 0.5 ) %		1381
$K \bar{K} \eta$	( 1.35 $\pm$ 0.16 ) %		1265
$\eta \pi^+ \pi^-$	( 1.7 $\pm$ 0.5 ) %		1427
$\eta 2(\pi^+ \pi^-)$	( 4.4 $\pm$ 1.3 ) %		1385
$K^+ K^- \pi^+ \pi^-$	( 6.9 $\pm$ 1.1 ) $\times 10^{-3}$		1345
$K^+ K^- \pi^+ \pi^- \pi^0$	( 3.5 $\pm$ 0.6 ) %		1304
$K^0 K^- \pi^+ \pi^- \pi^+ + c.c.$	( 5.6 $\pm$ 1.5 ) %		—
$K^+ K^- 2(\pi^+ \pi^-)$	( 7.5 $\pm$ 2.4 ) $\times 10^{-3}$		1253
$2(K^+ K^-)$	( 1.46 $\pm$ 0.30 ) $\times 10^{-3}$		1055
$\pi^+ \pi^- \pi^0 \pi^0$	( 4.7 $\pm$ 1.0 ) %		1460
$2(\pi^+ \pi^-)$	( 9.7 $\pm$ 1.2 ) $\times 10^{-3}$		1459
$2(\pi^+ \pi^- \pi^0)$	( 17.4 $\pm$ 3.3 ) %		1409
$3(\pi^+ \pi^-)$	( 1.8 $\pm$ 0.4 ) %		1406
$p \bar{p}$	( 1.50 $\pm$ 0.16 ) $\times 10^{-3}$		1160
$p \bar{p} \pi^0$	( 3.6 $\pm$ 1.3 ) $\times 10^{-3}$		1101
$\Lambda \bar{\Lambda}$	( 1.09 $\pm$ 0.24 ) $\times 10^{-3}$		990
$\Sigma^+ \bar{\Sigma}^-$	( 2.1 $\pm$ 0.6 ) $\times 10^{-3}$		900
$\Xi^- \bar{\Xi}^+$	( 8.9 $\pm$ 2.7 ) $\times 10^{-4}$		692
$\pi^+ \pi^- p \bar{p}$	( 5.3 $\pm$ 1.8 ) $\times 10^{-3}$		1027
<b>Radiative decays</b>			
$\gamma \gamma$	( 1.59 $\pm$ 0.13 ) $\times 10^{-4}$		1492
<b>Charge conjugation (C), Parity (P), Lepton family number (LF) violating modes</b>			
$\pi^+ \pi^-$	$P, CP < 1.1$ $\times 10^{-4}$	90%	1485
$\pi^0 \pi^0$	$P, CP < 4$ $\times 10^{-5}$	90%	1486
$K^+ K^-$	$P, CP < 6$ $\times 10^{-4}$	90%	1408
$K_S^0 K_S^0$	$P, CP < 3.1$ $\times 10^{-4}$	90%	1406

 **$J/\psi(1S)$** 

$$J^G(J^{PC}) = 0^-(1^{--})$$

Mass  $m = 3096.900 \pm 0.006$  MeV

Full width  $\Gamma = 92.9 \pm 2.8$  keV ( $S = 1.1$ )

$\Gamma_{ee} = 5.55 \pm 0.14 \pm 0.02$  keV

$J/\psi(1S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level (MeV/c)	$\rho$
hadrons	( 87.7 $\pm$ 0.5 ) %		—
virtual $\gamma \rightarrow$ hadrons	( 13.50 $\pm$ 0.30 ) %		—
$g g g$	( 64.1 $\pm$ 1.0 ) %		—
$\gamma g g$	( 8.8 $\pm$ 1.1 ) %		—
$e^+ e^-$	( 5.971 $\pm$ 0.032 ) %		1548
$e^+ e^- \gamma$	[ssaa] ( 8.8 $\pm$ 1.4 ) $\times 10^{-3}$		1548
$\mu^+ \mu^-$	( 5.961 $\pm$ 0.033 ) %		1545

**Decays involving hadronic resonances**

$\rho \pi$	( 1.69 $\pm$ 0.15 ) %	S=2.4	1448
$\rho^0 \pi^0$	( 5.6 $\pm$ 0.7 ) $\times 10^{-3}$		1448
$a_2(1320) \rho$	( 1.09 $\pm$ 0.22 ) %		1123
$\omega \pi^+ \pi^+ \pi^- \pi^-$	( 8.5 $\pm$ 3.4 ) $\times 10^{-3}$		1392
$\omega \pi^+ \pi^- \pi^0$	( 4.0 $\pm$ 0.7 ) $\times 10^{-3}$		1418
$\omega \pi^+ \pi^-$	( 8.6 $\pm$ 0.7 ) $\times 10^{-3}$	S=1.1	1435
$\omega f_2(1270)$	( 4.3 $\pm$ 0.6 ) $\times 10^{-3}$		1142
$K^*(892)^0 \bar{K}^*(892)^0$	( 2.3 $\pm$ 0.7 ) $\times 10^{-4}$		1266
$K^*(892)^\pm K^*(892)^\mp$	( 1.00 $\pm$ 0.22 $\pm$ 0.40 ) $\times 10^{-3}$		1266
$K^*(892)^\pm K^*(800)^\mp$	( 1.1 $\pm$ 1.0 $\pm$ 0.6 ) $\times 10^{-3}$		—
$\eta K^*(892)^0 \bar{K}^*(892)^0$	( 1.15 $\pm$ 0.26 ) $\times 10^{-3}$		1003
$K^*(892)^0 \bar{K}_2^*(1430)^0 + c.c.$	( 6.0 $\pm$ 0.6 ) $\times 10^{-3}$		1012
$K^*(892)^0 \bar{K}_2^*(1770)^0 + c.c. \rightarrow$	( 6.9 $\pm$ 0.9 ) $\times 10^{-4}$		—
$K^*(892)^0 \bar{K}^- \pi^+ + c.c.$			—
$\omega K^*(892) \bar{K} + c.c.$	( 6.1 $\pm$ 0.9 ) $\times 10^{-3}$		1097
$K^+ K^*(892)^- + c.c.$	( 5.12 $\pm$ 0.30 ) $\times 10^{-3}$		1373
$K^+ K^*(892)^- + c.c. \rightarrow$	( 1.97 $\pm$ 0.20 ) $\times 10^{-3}$		—
$K^+ K^- \pi^0$			—
$K^+ K^*(892)^- + c.c. \rightarrow$	( 3.0 $\pm$ 0.4 ) $\times 10^{-3}$		—
$K^0 K^\pm \pi^\mp + c.c.$			—
$K^0 \bar{K}^*(892)^0 + c.c.$	( 4.39 $\pm$ 0.31 ) $\times 10^{-3}$		1373
$K^0 \bar{K}^*(892)^0 + c.c. \rightarrow$	( 3.2 $\pm$ 0.4 ) $\times 10^{-3}$		—
$K^0 K^\pm \pi^\mp + c.c.$			—
$K_1(1400)^\pm K^\mp$	( 3.8 $\pm$ 1.4 ) $\times 10^{-3}$		1170
$\bar{K}^*(892)^0 K^+ \pi^- + c.c.$	seen		1343
$\omega \pi^0 \pi^0$	( 3.4 $\pm$ 0.8 ) $\times 10^{-3}$		1436
$b_1(1235)^\pm \pi^\mp$	[hh] ( 3.0 $\pm$ 0.5 ) $\times 10^{-3}$		1300
$\omega K^\pm K_S^0 \pi^\mp$	[hh] ( 3.4 $\pm$ 0.5 ) $\times 10^{-3}$		1210
$b_1(1235)^0 \pi^0$	( 2.3 $\pm$ 0.6 ) $\times 10^{-3}$		1300
$\eta K^\pm K_S^0 \pi^\mp$	[hh] ( 2.2 $\pm$ 0.4 ) $\times 10^{-3}$		1278
$\phi K^*(892) \bar{K} + c.c.$	( 2.18 $\pm$ 0.23 ) $\times 10^{-3}$		969
$\omega K \bar{K}$	( 1.70 $\pm$ 0.32 ) $\times 10^{-3}$		1268
$\omega f_0(1710) \rightarrow \omega K \bar{K}$	( 4.8 $\pm$ 1.1 ) $\times 10^{-4}$		878
$\phi 2(\pi^+ \pi^-)$	( 1.66 $\pm$ 0.23 ) $\times 10^{-3}$		1318
$\Delta(1232)^{++} \bar{p} \pi^-$	( 1.6 $\pm$ 0.5 ) $\times 10^{-3}$		1030
$\omega \eta$	( 1.74 $\pm$ 0.20 ) $\times 10^{-3}$	S=1.6	1394
$\phi K \bar{K}$	( 1.83 $\pm$ 0.24 ) $\times 10^{-3}$	S=1.5	1179
$\phi f_0(1710) \rightarrow \phi K \bar{K}$	( 3.6 $\pm$ 0.6 ) $\times 10^{-4}$		875
$\phi f_2(1270)$	( 7.2 $\pm$ 1.3 ) $\times 10^{-4}$		1036
$\Delta(1232)^{++} \bar{\Delta}(1232)^{--}$	( 1.10 $\pm$ 0.29 ) $\times 10^{-3}$		938
$\Sigma(1385)^- \bar{\Sigma}(1385)^+ (or c.c.)$	[hh] ( 1.10 $\pm$ 0.12 ) $\times 10^{-3}$		697
$\phi f_2'(1525)$	( 8 $\pm$ 4 ) $\times 10^{-4}$	S=2.7	871
$\phi \pi^+ \pi^-$	( 9.4 $\pm$ 0.9 ) $\times 10^{-4}$	S=1.2	1365
$\phi \pi^0 \pi^0$	( 5.6 $\pm$ 1.6 ) $\times 10^{-4}$		1366
$\phi K^\pm K_S^0 \pi^\mp$	[hh] ( 7.2 $\pm$ 0.8 ) $\times 10^{-4}$		1114
$\omega f_1(1420)$	( 6.8 $\pm$ 2.4 ) $\times 10^{-4}$		1062
$\phi \eta$	( 7.5 $\pm$ 0.8 ) $\times 10^{-4}$	S=1.5	1320
$\Xi^0 \Xi^0$	( 1.20 $\pm$ 0.24 ) $\times 10^{-3}$		818
$\Xi(1530)^- \Xi^+$	( 5.9 $\pm$ 1.5 ) $\times 10^{-4}$		600
$p K^- \bar{\Sigma}(1385)^0$	( 5.1 $\pm$ 3.2 ) $\times 10^{-4}$		646
$\omega \pi^0$	( 4.5 $\pm$ 0.5 ) $\times 10^{-4}$	S=1.4	1446
$\phi \eta'(958)$	( 4.0 $\pm$ 0.7 ) $\times 10^{-4}$	S=2.1	1192
$\phi f_0(980)$	( 3.2 $\pm$ 0.9 ) $\times 10^{-4}$	S=1.9	1178
$\phi f_0(980) \rightarrow \phi \pi^+ \pi^-$	( 1.8 $\pm$ 0.4 ) $\times 10^{-4}$		—
$\phi f_0(980) \rightarrow \phi \pi^0 \pi^0$	( 1.7 $\pm$ 0.7 ) $\times 10^{-4}$		—
$\phi \pi^0 f_0(980) \rightarrow \phi \pi^0 \pi^+ \pi^-$	( 4.5 $\pm$ 1.0 ) $\times 10^{-6}$		—
$\phi \pi^0 f_0(980) \rightarrow \phi \pi^0 p^0 \pi^0$	( 1.7 $\pm$ 0.6 ) $\times 10^{-6}$		1045
$\eta \phi f_0(980) \rightarrow \eta \phi \pi^+ \pi^-$	( 3.2 $\pm$ 1.0 ) $\times 10^{-4}$		—
$\phi a_0(980)^0 \rightarrow \phi \eta \pi^0$	( 5 $\pm$ 4 ) $\times 10^{-6}$		—
$\Xi(1530)^0 \Xi^0$	( 3.2 $\pm$ 1.4 ) $\times 10^{-4}$		608

# Meson Summary Table

$\Sigma(1385)^-\bar{\Sigma}^+$ (or c.c.)	[hh]	( 3.1 ± 0.5 ) × 10 <sup>-4</sup>		855	$\rho K^-\bar{\Sigma}^0$	( 2.9 ± 0.8 ) × 10 <sup>-4</sup>	819
$\phi f_1(1285)$		( 2.6 ± 0.5 ) × 10 <sup>-4</sup>		1032	$K^+ K^-$	( 2.86 ± 0.21 ) × 10 <sup>-4</sup>	1468
$\phi f_1(1285) \rightarrow$		( 9.4 ± 2.8 ) × 10 <sup>-7</sup>		952	$K_S^0 K_L^0$	( 2.1 ± 0.4 ) × 10 <sup>-4</sup>	S=3.2 1466
$\phi \pi^0 f_0(980) \rightarrow$					$\Lambda \bar{\Lambda} \pi^+ \pi^-$	( 4.3 ± 1.0 ) × 10 <sup>-3</sup>	903
$\phi \pi^0 \pi^+ \pi^-$					$\Lambda \bar{\Lambda} \eta$	( 1.62 ± 0.17 ) × 10 <sup>-4</sup>	672
$\phi f_1(1285) \rightarrow$		( 2.1 ± 2.2 ) × 10 <sup>-7</sup>		955	$\Lambda \bar{\Lambda} \pi^0$	( 3.8 ± 0.4 ) × 10 <sup>-5</sup>	998
$\phi \pi^0 f_0(980) \rightarrow$					$\bar{\Lambda} n K_S^0 + \text{c.c.}$	( 6.5 ± 1.1 ) × 10 <sup>-4</sup>	872
$\phi \pi^0 \pi^0 \pi^0$					$\pi^+ \pi^-$	( 1.47 ± 0.14 ) × 10 <sup>-4</sup>	1542
$\eta \pi^+ \pi^-$		( 4.0 ± 1.7 ) × 10 <sup>-4</sup>		1487	$\Lambda \bar{\Sigma}^+ + \text{c.c.}$	( 2.83 ± 0.23 ) × 10 <sup>-5</sup>	1034
$\eta \rho$		( 1.93 ± 0.23 ) × 10 <sup>-4</sup>		1396	$K_S^0 K_S^0$	< 1 × 10 <sup>-6</sup>	CL=95% 1466
$\omega \eta'(958)$		( 1.82 ± 0.21 ) × 10 <sup>-4</sup>		1279	Radiative decays		
$\omega f_0(980)$		( 1.4 ± 0.5 ) × 10 <sup>-4</sup>		1267	$3\gamma$	( 1.16 ± 0.22 ) × 10 <sup>-5</sup>	1548
$\rho \eta'(958)$		( 1.05 ± 0.18 ) × 10 <sup>-4</sup>		1281	$4\gamma$	< 9 × 10 <sup>-6</sup>	CL=90% 1548
$a_2(1320)^{\pm} \pi^{\mp}$	[hh]	< 4.3 × 10 <sup>-3</sup>	CL=90%	1263	$5\gamma$	< 1.5 × 10 <sup>-5</sup>	CL=90% 1548
$K \bar{K}_2^*(1430) + \text{c.c.}$		< 4.0 × 10 <sup>-3</sup>	CL=90%	1159	$\gamma \pi^0 \pi^0$	( 1.15 ± 0.05 ) × 10 <sup>-3</sup>	1543
$K_1(1270)^{\pm} K^{\mp}$		< 3.0 × 10 <sup>-3</sup>	CL=90%	1231	$\gamma \eta_c(1S)$	( 1.7 ± 0.4 ) %	S=1.5 111
$K_2^*(1430)^0 \bar{K}_2^*(1430)^0$		< 2.9 × 10 <sup>-3</sup>	CL=90%	604	$\gamma \eta_c(1S) \rightarrow 3\gamma$	( 3.8 ± 1.3 -1.0 ) × 10 <sup>-6</sup>	S=1.1 -
$\phi \pi^0$		3 × 10 <sup>-6</sup> or 1 × 10 <sup>-7</sup>		1377	$\gamma \pi^+ \pi^- 2\pi^0$	( 8.3 ± 3.1 ) × 10 <sup>-3</sup>	1518
$\phi \eta(1405) \rightarrow \phi \eta \pi^+ \pi^-$		( 2.0 ± 1.0 ) × 10 <sup>-5</sup>		946	$\gamma \eta \pi \pi$	( 6.1 ± 1.0 ) × 10 <sup>-3</sup>	1487
$\omega f_2'(1525)$		< 2.2 × 10 <sup>-4</sup>	CL=90%	1003	$\gamma \eta_2(1870) \rightarrow \gamma \eta \pi^+ \pi^-$	( 6.2 ± 2.4 ) × 10 <sup>-4</sup>	-
$\omega X(1835) \rightarrow \omega \rho \bar{\rho}$		< 3.9 × 10 <sup>-6</sup>	CL=95%	-	$\gamma \eta(1405/1475) \rightarrow \gamma K \bar{K} \pi$	[o] ( 2.8 ± 0.6 ) × 10 <sup>-3</sup>	S=1.6 1223
$\phi X(1835) \rightarrow \phi \eta \pi^+ \pi^-$		< 2.8 × 10 <sup>-4</sup>	CL=90%	578	$\gamma \eta(1405/1475) \rightarrow \gamma \gamma \rho^0$	( 7.8 ± 2.0 ) × 10 <sup>-5</sup>	S=1.8 1223
$\phi X(1870) \rightarrow \phi \eta \pi^+ \pi^-$		< 6.13 × 10 <sup>-5</sup>	CL=90%	-	$\gamma \eta(1405/1475) \rightarrow \gamma \eta \pi^+ \pi^-$	( 3.0 ± 0.5 ) × 10 <sup>-4</sup>	-
$\eta \phi(2170) \rightarrow \eta \phi f_0(980) \rightarrow$		( 1.2 ± 0.4 ) × 10 <sup>-4</sup>		628	$\gamma \eta(1405/1475) \rightarrow \gamma \gamma \phi$	< 8.2 × 10 <sup>-5</sup>	CL=95% -
$\eta \phi \pi^+ \pi^-$					$\gamma \rho \rho$	( 4.5 ± 0.8 ) × 10 <sup>-3</sup>	1340
$\eta \phi(2170) \rightarrow$		< 2.52 × 10 <sup>-4</sup>	CL=90%	-	$\gamma \rho \omega$	< 5.4 × 10 <sup>-4</sup>	CL=90% 1338
$\eta K^*(892)^0 \bar{K}^*(892)^0$					$\gamma \rho \phi$	< 8.8 × 10 <sup>-5</sup>	CL=90% 1258
$\Sigma(1385)^0 \bar{\Lambda} + \text{c.c.}$		< 8.2 × 10 <sup>-6</sup>	CL=90%	912	$\gamma \eta'(958)$	( 5.15 ± 0.16 ) × 10 <sup>-3</sup>	S=1.2 1400
$\Delta(1232)^+ \bar{p}$		< 1 × 10 <sup>-4</sup>	CL=90%	1100	$\gamma 2\pi^+ 2\pi^-$	( 2.8 ± 0.5 ) × 10 <sup>-3</sup>	S=1.9 1517
$\Lambda(1520) \bar{\Lambda} + \text{c.c.} \rightarrow \gamma \Lambda \bar{\Lambda}$		< 4.1 × 10 <sup>-6</sup>	CL=90%	-	$\gamma f_2(1270) f_2(1270)$	( 9.5 ± 1.7 ) × 10 <sup>-4</sup>	878
$\Theta(1540) \bar{\Theta}(1540) \rightarrow$		< 1.1 × 10 <sup>-5</sup>	CL=90%	-	$\gamma f_2(1270) f_2(1270) (\text{non reso-})$	( 8.2 ± 1.9 ) × 10 <sup>-4</sup>	-
$K_S^0 p K^- \bar{n} + \text{c.c.}$					$\gamma K^+ K^- \pi^+ \pi^-$	( 2.1 ± 0.6 ) × 10 <sup>-3</sup>	1407
$\Theta(1540) K^- \bar{n} \rightarrow K_S^0 p K^- \bar{n}$		< 2.1 × 10 <sup>-5</sup>	CL=90%	-	$\gamma f_4(2050)$	( 2.7 ± 0.7 ) × 10 <sup>-3</sup>	891
$\Theta(1540) K_S^0 \bar{p} \rightarrow K_S^0 \bar{p} K^+ n$		< 1.6 × 10 <sup>-5</sup>	CL=90%	-	$\gamma \omega \omega$	( 1.61 ± 0.33 ) × 10 <sup>-3</sup>	1336
$\bar{\Theta}(1540) K^+ n \rightarrow K_S^0 \bar{p} K^+ n$		< 5.6 × 10 <sup>-5</sup>	CL=90%	-	$\gamma \eta(1405/1475) \rightarrow \gamma \rho^0 \rho^0$	( 1.7 ± 0.4 ) × 10 <sup>-3</sup>	S=1.3 1223
$\bar{\Theta}(1540) K_S^0 p \rightarrow K_S^0 p K^- \bar{n}$		< 1.1 × 10 <sup>-5</sup>	CL=90%	-	$\gamma f_2(1270)$	( 1.64 ± 0.12 ) × 10 <sup>-3</sup>	S=1.3 1286
$\Sigma^0 \bar{\Lambda}$		< 9 × 10 <sup>-5</sup>	CL=90%	1032	$\gamma f_0(1370) \rightarrow \gamma K \bar{K}$	( 4.2 ± 1.5 ) × 10 <sup>-4</sup>	-
Decays into stable hadrons					$\gamma f_0(1710) \rightarrow \gamma K \bar{K}$	( 1.00 ± 0.11 -0.09 ) × 10 <sup>-3</sup>	S=1.5 1075
$2(\pi^+ \pi^-) \pi^0$		( 4.1 ± 0.5 ) %	S=2.4	1496	$\gamma f_0(1710) \rightarrow \gamma \pi \pi$	( 3.8 ± 0.5 ) × 10 <sup>-4</sup>	-
$3(\pi^+ \pi^-) \pi^0$		( 2.9 ± 0.6 ) %		1433	$\gamma f_0(1710) \rightarrow \gamma \omega \omega$	( 3.1 ± 1.0 ) × 10 <sup>-4</sup>	-
$\pi^+ \pi^- \pi^0$		( 2.11 ± 0.07 ) %	S=1.5	1533	$\gamma f_0(1710) \rightarrow \gamma \eta \eta$	( 2.4 ± 1.2 -0.7 ) × 10 <sup>-4</sup>	-
$\pi^+ \pi^- \pi^0 K^+ K^-$		( 1.79 ± 0.29 ) %	S=2.2	1368	$\gamma \eta$	( 1.104 ± 0.034 ) × 10 <sup>-3</sup>	1500
$4(\pi^+ \pi^-) \pi^0$		( 9.0 ± 3.0 ) × 10 <sup>-3</sup>		1345	$\gamma f_1(1420) \rightarrow \gamma K \bar{K} \pi$	( 7.9 ± 1.3 ) × 10 <sup>-4</sup>	1220
$\pi^+ \pi^- K^+ K^-$		( 6.6 ± 0.5 ) × 10 <sup>-3</sup>		1407	$\gamma f_1(1285)$	( 6.1 ± 0.8 ) × 10 <sup>-4</sup>	1283
$\pi^+ \pi^- K^+ K^- \eta$		( 1.84 ± 0.28 ) × 10 <sup>-3</sup>		1221	$\gamma f_1(1510) \rightarrow \gamma \eta \pi^+ \pi^-$	( 4.5 ± 1.2 ) × 10 <sup>-4</sup>	-
$\pi^0 \pi^0 K^+ K^-$		( 2.45 ± 0.31 ) × 10 <sup>-3</sup>		1410	$\gamma f_2'(1525)$	( 5.7 ± 0.8 -0.5 ) × 10 <sup>-4</sup>	S=1.5 1173
$K \bar{K} \pi$		( 6.1 ± 1.0 ) × 10 <sup>-3</sup>		1442	$\gamma f_2'(1525) \rightarrow \gamma \eta \eta$	( 3.4 ± 1.4 ) × 10 <sup>-5</sup>	-
$2(\pi^+ \pi^-)$		( 3.57 ± 0.30 ) × 10 <sup>-3</sup>		1517	$\gamma f_2(1640) \rightarrow \gamma \omega \omega$	( 2.8 ± 1.8 ) × 10 <sup>-4</sup>	-
$3(\pi^+ \pi^-)$		( 4.3 ± 0.4 ) × 10 <sup>-3</sup>		1466	$\gamma f_2(1910) \rightarrow \gamma \omega \omega$	( 2.0 ± 1.4 ) × 10 <sup>-4</sup>	-
$2(\pi^+ \pi^- \pi^0)$		( 1.62 ± 0.21 ) %		1468	$\gamma f_0(1800) \rightarrow \gamma \omega \phi$	( 2.5 ± 0.6 ) × 10 <sup>-4</sup>	-
$2(\pi^+ \pi^-) \eta$		( 2.29 ± 0.24 ) × 10 <sup>-3</sup>		1446	$\gamma f_2(1810) \rightarrow \gamma \eta \eta$	( 5.4 ± 3.5 -2.4 ) × 10 <sup>-5</sup>	-
$3(\pi^+ \pi^-) \eta$		( 7.2 ± 1.5 ) × 10 <sup>-4</sup>		1379	$\gamma f_2(1950) \rightarrow$	( 7.0 ± 2.2 ) × 10 <sup>-4</sup>	-
$\rho \bar{\rho}$		( 2.120 ± 0.029 ) × 10 <sup>-3</sup>		1232	$\gamma K^*(892) \bar{K}^*(892)$		
$\rho \bar{\rho} \pi^0$		( 1.19 ± 0.08 ) × 10 <sup>-3</sup>	S=1.1	1176	$\gamma K^*(892) \bar{K}^*(892)$	( 4.0 ± 1.3 ) × 10 <sup>-3</sup>	1266
$\rho \bar{\rho} \pi^+ \pi^-$		( 6.0 ± 0.5 ) × 10 <sup>-3</sup>	S=1.3	1107	$\gamma \phi \phi$	( 4.0 ± 1.2 ) × 10 <sup>-4</sup>	S=2.1 1166
$\rho \bar{\rho} \pi^+ \pi^- \pi^0$	[ttaa]	( 2.3 ± 0.9 ) × 10 <sup>-3</sup>	S=1.9	1033	$\gamma \rho \bar{\rho}$	( 3.8 ± 1.0 ) × 10 <sup>-4</sup>	1232
$\rho \bar{\rho} \eta$		( 2.00 ± 0.12 ) × 10 <sup>-3</sup>		948	$\gamma \eta(2225)$	( 3.3 ± 0.5 ) × 10 <sup>-4</sup>	749
$\rho \bar{\rho} \rho$		< 3.1 × 10 <sup>-4</sup>	CL=90%	774	$\gamma \eta(1760) \rightarrow \gamma \rho^0 \rho^0$	( 1.3 ± 0.9 ) × 10 <sup>-4</sup>	1048
$\rho \bar{\rho} \omega$		( 9.8 ± 1.0 ) × 10 <sup>-4</sup>	S=1.3	768	$\gamma \eta(1760) \rightarrow \gamma \omega \omega$	( 1.98 ± 0.33 ) × 10 <sup>-3</sup>	-
$\rho \bar{\rho} \eta'(958)$		( 2.1 ± 0.4 ) × 10 <sup>-4</sup>		596	$\gamma X(1835) \rightarrow \gamma \pi^+ \pi^- \eta'$	( 2.6 ± 0.4 ) × 10 <sup>-4</sup>	1006
$\rho \bar{\rho} a_0(980) \rightarrow \rho \bar{\rho} \pi^0 \eta$		( 6.8 ± 1.8 ) × 10 <sup>-5</sup>		-	$\gamma X(1835) \rightarrow \gamma \rho \bar{\rho}$	( 7.7 ± 1.5 -0.9 ) × 10 <sup>-5</sup>	-
$\rho \bar{\rho} \phi$		( 4.5 ± 1.5 ) × 10 <sup>-5</sup>		527	$\gamma X(1835) \rightarrow \gamma K_S^0 K_S^0 \eta$	( 3.3 ± 2.0 -1.3 ) × 10 <sup>-5</sup>	-
$n \bar{n}$		( 2.09 ± 0.16 ) × 10 <sup>-3</sup>		1231	$\gamma X(1840) \rightarrow \gamma 3(\pi^+ \pi^-)$	( 2.4 ± 0.7 -0.8 ) × 10 <sup>-5</sup>	-
$n \bar{n} \pi^+ \pi^-$		( 4 ± 4 ) × 10 <sup>-3</sup>		1106	$\gamma (K \bar{K} \pi) [J^{PC} = 0^- +]$	( 7 ± 4 ) × 10 <sup>-4</sup>	S=2.1 1442
$\Sigma^+ \bar{\Sigma}^-$		( 1.50 ± 0.24 ) × 10 <sup>-3</sup>		992	$\gamma \pi^0$	( 3.49 ± 0.33 -0.30 ) × 10 <sup>-5</sup>	1546
$\Sigma^0 \bar{\Sigma}^0$		( 1.29 ± 0.09 ) × 10 <sup>-3</sup>		988	$\gamma \rho \bar{\rho} \pi^+ \pi^-$	< 7.9 × 10 <sup>-4</sup>	CL=90% 1107
$2(\pi^+ \pi^-) K^+ K^-$		( 4.7 ± 0.7 ) × 10 <sup>-3</sup>	S=1.3	1320	$\gamma \Lambda \bar{\Lambda}$	< 1.3 × 10 <sup>-4</sup>	CL=90% 1074
$\rho \bar{n} \pi^-$		( 2.12 ± 0.09 ) × 10 <sup>-3</sup>		1174	$\gamma f_0(2100) \rightarrow \gamma \eta \eta$	( 1.13 ± 0.60 -0.30 ) × 10 <sup>-4</sup>	-
$n N(1440)$	seen			984	$\gamma f_0(2100) \rightarrow \gamma \pi \pi$	( 6.2 ± 1.0 ) × 10 <sup>-4</sup>	-
$n N(1520)$	seen			928			
$n N(1535)$	seen			914			
$\Xi^- \bar{\Xi}^+$		( 8.6 ± 1.1 ) × 10 <sup>-4</sup>	S=1.2	807			
$\Lambda \bar{\Lambda}$		( 1.61 ± 0.15 ) × 10 <sup>-3</sup>	S=1.9	1074			
$\Lambda \bar{\Sigma}^- \pi^+$ (or c.c.)	[hh]	( 8.3 ± 0.7 ) × 10 <sup>-4</sup>	S=1.2	950			
$\rho K^- \bar{\Lambda}$		( 8.9 ± 1.6 ) × 10 <sup>-4</sup>		876			
$2(K^+ K^-)$		( 7.6 ± 0.9 ) × 10 <sup>-4</sup>		1131			



## Meson Summary Table

$\gamma f_0(2200) \rightarrow \gamma K \bar{K}$	$(5.9 \pm 1.3) \times 10^{-4}$		$K_S^0 K_S^0 \pi^+ \pi^-$	$(5.6 \pm 1.0) \times 10^{-3}$	1579			
$\gamma f_J(2220) \rightarrow \gamma \pi \pi$	$< 3.9 \times 10^{-5}$	CL=90%	$K^+ K^- \eta \pi^0$	$(3.0 \pm 0.7) \times 10^{-3}$	1468			
$\gamma f_J(2220) \rightarrow \gamma K \bar{K}$	$< 4.1 \times 10^{-5}$	CL=90%	$3(\pi^+ \pi^-)$	$(1.20 \pm 0.18) \%$	1633			
$\gamma f_J(2220) \rightarrow \gamma p \bar{p}$	$(1.5 \pm 0.8) \times 10^{-5}$		$K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.}$	$(7.2 \pm 1.6) \times 10^{-3}$	1523			
$\gamma f_2(2340) \rightarrow \gamma \eta \eta$	$(5.6 \pm_{-2.2}^{+2.4}) \times 10^{-5}$		$K^*(892)^0 \bar{K}^*(892)^0$	$(1.7 \pm 0.6) \times 10^{-3}$	1456			
$\gamma f_0(1500) \rightarrow \gamma \pi \pi$	$(1.09 \pm 0.24) \times 10^{-4}$	1183	$\pi \pi$	$(8.33 \pm 0.35) \times 10^{-3}$	1702			
$\gamma f_0(1500) \rightarrow \gamma \eta \eta$	$(1.7 \pm_{-1.4}^{+0.6}) \times 10^{-5}$		$\pi^0 \eta$	$< 1.8 \times 10^{-4}$	1661			
$\gamma A \rightarrow \gamma \text{invisible}$	$[uu\bar{a}] < 6.3 \times 10^{-6}$	CL=90%	$\pi^0 \eta'$	$< 1.1 \times 10^{-3}$	1570			
$\gamma A^0 \rightarrow \gamma \mu^+ \mu^-$	$[v\bar{v}a\bar{a}] < 2.1 \times 10^{-5}$	CL=90%	$\pi^0 \eta_c$	$< 1.6 \times 10^{-3}$	CL=90% 384			
<b>Dalitz decays</b>			$\eta \eta$	$(2.95 \pm 0.19) \times 10^{-3}$	1617			
$\pi^0 e^+ e^-$	$(7.6 \pm 1.4) \times 10^{-7}$	1546	$\eta \eta'$	$< 2.3 \times 10^{-4}$	CL=90% 1521			
$\eta e^+ e^-$	$(1.16 \pm 0.09) \times 10^{-5}$	1500	$\eta' \eta'$	$(1.96 \pm 0.21) \times 10^{-3}$	1413			
$\eta'(958) e^+ e^-$	$(5.81 \pm 0.35) \times 10^{-5}$	1400	$\omega \omega$	$(9.5 \pm 1.1) \times 10^{-4}$	1517			
<b>Weak decays</b>			$\omega \phi$	$(1.16 \pm 0.21) \times 10^{-4}$	1447			
$D^- e^+ \nu_e + \text{c.c.}$	$< 1.2 \times 10^{-5}$	CL=90% 984	$\omega K^+ K^-$	$(1.94 \pm 0.21) \times 10^{-3}$	1457			
$\bar{D}^0 e^+ e^- + \text{c.c.}$	$< 1.1 \times 10^{-5}$	CL=90% 987	$K^+ K^-$	$(5.91 \pm 0.32) \times 10^{-3}$	1634			
$D_s^- e^+ \nu_e + \text{c.c.}$	$< 1.3 \times 10^{-6}$	CL=90% 923	$K_S^0 K_S^0$	$(3.10 \pm 0.18) \times 10^{-3}$	1633			
$D_s^{*-} e^+ \nu_e + \text{c.c.}$	$< 1.8 \times 10^{-6}$	CL=90% 828	$\pi^+ \pi^- \eta$	$< 1.9 \times 10^{-4}$	CL=90% 1651			
$D^- \pi^+ + \text{c.c.}$	$< 7.5 \times 10^{-5}$	CL=90% 977	$\pi^+ \pi^- \eta'$	$< 3.5 \times 10^{-4}$	CL=90% 1560			
$\bar{D}^0 K^0 + \text{c.c.}$	$< 1.7 \times 10^{-4}$	CL=90% 898	$\bar{K}^0 K^+ \pi^- + \text{c.c.}$	$< 9 \times 10^{-5}$	CL=90% 1610			
$\bar{D}^0 \bar{K}^{*0} + \text{c.c.}$	$< 2.5 \times 10^{-6}$	CL=90% 670	$K^+ K^- \pi^0$	$< 6 \times 10^{-5}$	CL=90% 1611			
$D_s^- \pi^+ + \text{c.c.}$	$< 1.3 \times 10^{-4}$	CL=90% 916	$K^+ K^- \eta$	$< 2.2 \times 10^{-4}$	CL=90% 1512			
$D_s^- \rho^+ + \text{c.c.}$	$< 1.3 \times 10^{-5}$	CL=90% 663	$K^+ K^- K_S^0 K_S^0$	$(1.4 \pm 0.5) \times 10^{-3}$	1331			
<b>Charge conjugation (C), Parity (P), Lepton Family number (LF) violating modes</b>			$K^+ K^- K^+ K^-$	$(2.75 \pm 0.28) \times 10^{-3}$	1333			
$\gamma \gamma$	C	$< 2.7 \times 10^{-7}$	CL=90% 1548	$K^+ K^- \phi$	$(9.5 \pm 2.4) \times 10^{-4}$	1381		
$\gamma \phi$	C	$< 1.4 \times 10^{-6}$	CL=90% 1381	$\bar{K}^0 K^+ \pi^- \phi + \text{c.c.}$	$(3.7 \pm 0.6) \times 10^{-3}$	1326		
$e^\pm \mu^\mp$	LF	$< 1.6 \times 10^{-7}$	CL=90% 1547	$K^+ K^- \pi^0 \phi$	$(1.90 \pm 0.35) \times 10^{-3}$	1329		
$e^\pm \tau^\mp$	LF	$< 8.3 \times 10^{-6}$	CL=90% 1039	$\phi \pi^+ \pi^- \pi^0$	$(1.18 \pm 0.15) \times 10^{-3}$	1525		
$\mu^\pm \tau^\mp$	LF	$< 2.0 \times 10^{-6}$	CL=90% 1035	$\phi \phi$	$(7.7 \pm 0.7) \times 10^{-4}$	1370		
<b>Other decays</b>			$p \bar{p}$	$(2.25 \pm 0.09) \times 10^{-4}$	1426			
invisible	$< 7 \times 10^{-4}$	CL=90%	-	$p \bar{p} \pi^0$	$(6.8 \pm 0.7) \times 10^{-4}$	S=1.3 1379		
<div><div><math>\chi_{c0}(1P)</math></div><div><math>I^G(J^{PC}) = 0^+(0^{++})</math></div><div>Mass <math>m = 3414.75 \pm 0.31</math> MeV</div><div>Full width <math>\Gamma = 10.5 \pm 0.6</math> MeV</div></div>						$p \bar{p} \eta$	$(3.5 \pm 0.4) \times 10^{-4}$	1187
<div><div><math>\chi_{c0}(1P)</math> DECAY MODES</div><div>Fraction (<math>\Gamma_i/\Gamma</math>)</div><div>Scale factor/ Confidence level</div><div><math>p</math> (MeV/c)</div></div>						$p \bar{p} \omega$	$(5.1 \pm 0.6) \times 10^{-4}$	1043
<b>Hadronic decays</b>						$p \bar{p} \phi$	$(5.9 \pm 1.4) \times 10^{-5}$	876
$2(\pi^+ \pi^-)$	$(2.24 \pm 0.18) \%$		1679	$p \bar{p} \pi^+ \pi^-$	$(2.1 \pm 0.7) \times 10^{-3}$	S=1.4 1320		
$\rho^0 \pi^+ \pi^-$	$(8.7 \pm 2.8) \times 10^{-3}$		1607	$p \bar{p} \pi^0 \pi^0$	$(1.02 \pm 0.27) \times 10^{-3}$	1324		
$f_0(980) f_0(980)$	$(6.5 \pm 2.1) \times 10^{-4}$		1391	$p \bar{p} K^+ K^-$ (non-resonant)	$(1.19 \pm 0.26) \times 10^{-4}$	890		
$\pi^+ \pi^- \pi^0 \pi^0$	$(3.3 \pm 0.4) \%$		1680	$p \bar{p} K_S^0 K_S^0$	$< 8.8 \times 10^{-4}$	CL=90% 884		
$\rho^+ \pi^- \pi^0 + \text{c.c.}$	$(2.8 \pm 0.4) \%$		1607	$p \bar{n} \pi^-$	$(1.24 \pm 0.11) \times 10^{-3}$	1376		
$4\pi^0$	$(3.2 \pm 0.4) \times 10^{-3}$		1681	$\bar{p} n \pi^+$	$(1.34 \pm 0.12) \times 10^{-3}$	1376		
$\pi^+ \pi^- K^+ K^-$	$(1.75 \pm 0.14) \%$		1580	$p \bar{n} \pi^- \pi^0$	$(2.29 \pm 0.21) \times 10^{-3}$	1321		
$K_0^*(1430)^0 \bar{K}_0^*(1430)^0 \rightarrow$	$(9.6 \pm_{-2.8}^{+3.5}) \times 10^{-4}$	-		$\bar{p} n \pi^+ \pi^0$	$(2.16 \pm 0.18) \times 10^{-3}$	1321		
$\pi^+ \pi^- K^+ K^-$				$\Lambda \bar{\Lambda}$	$(3.21 \pm 0.25) \times 10^{-4}$	1292		
$K_0^*(1430)^0 \bar{K}_2^*(1430)^0 + \text{c.c.} \rightarrow$	$(7.8 \pm_{-2.4}^{+1.9}) \times 10^{-4}$	-		$\Lambda \bar{\Lambda} \pi^+ \pi^-$	$(1.15 \pm 0.13) \times 10^{-3}$	1153		
$\pi^+ \pi^- K^+ K^-$				$\Lambda \bar{\Lambda} \pi^+ \pi^-$ (non-resonant)	$< 5 \times 10^{-4}$	CL=90% 1153		
$K_1(1270)^+ K^- + \text{c.c.} \rightarrow$	$(6.1 \pm 1.9) \times 10^{-3}$	-		$\Sigma(1385)^+ \bar{\Lambda} \pi^- + \text{c.c.}$	$< 5 \times 10^{-4}$	CL=90% 1083		
$\pi^+ \pi^- K^+ K^-$				$\Sigma(1385)^- \bar{\Lambda} \pi^+ + \text{c.c.}$	$< 5 \times 10^{-4}$	CL=90% 1083		
$K_1(1400)^+ K^- + \text{c.c.} \rightarrow$	$< 2.6 \times 10^{-3}$	CL=90%	-	$K^+ \bar{p} \Lambda + \text{c.c.}$	$(1.22 \pm 0.12) \times 10^{-3}$	S=1.3 1132		
$\pi^+ \pi^- K^+ K^-$				$K^+ \bar{p} \Lambda(1520) + \text{c.c.}$	$(2.9 \pm 0.7) \times 10^{-4}$	858		
$f_0(980) f_0(980)$	$(1.6 \pm_{-0.9}^{+1.0}) \times 10^{-4}$		1391	$\Lambda(1520) \bar{\Lambda}(1520)$	$(3.1 \pm 1.2) \times 10^{-4}$	779		
$f_0(980) f_0(2200)$	$(7.8 \pm_{-2.5}^{+2.0}) \times 10^{-4}$		584	$\Sigma^0 \bar{\Sigma}^0$	$(4.4 \pm 0.4) \times 10^{-4}$	1222		
$f_0(1370) f_0(1370)$	$< 2.7 \times 10^{-4}$	CL=90%	1019	$\Sigma^+ \bar{\Sigma}^-$	$(3.9 \pm 0.7) \times 10^{-4}$	S=1.7 1225		
$f_0(1370) f_0(1500)$	$< 1.7 \times 10^{-4}$	CL=90%	921	$\Sigma(1385)^+ \bar{\Sigma}(1385)^-$	$(1.6 \pm 0.6) \times 10^{-4}$	1001		
$f_0(1370) f_0(1710)$	$(6.6 \pm_{-2.3}^{+3.5}) \times 10^{-4}$		720	$\Sigma(1385)^- \bar{\Sigma}(1385)^+$	$(2.3 \pm 0.6) \times 10^{-4}$	1001		
$f_0(1500) f_0(1370)$	$< 1.3 \times 10^{-4}$	CL=90%	921	$K^- \Lambda \bar{\Xi}^{++} + \text{c.c.}$	$(1.90 \pm 0.34) \times 10^{-4}$	873		
$f_0(1500) f_0(1500)$	$< 5 \times 10^{-5}$	CL=90%	807	$\Xi^0 \bar{\Xi}^0$	$(3.1 \pm 0.8) \times 10^{-4}$	1089		
$f_0(1500) f_0(1710)$	$< 7 \times 10^{-5}$	CL=90%	557	$\Xi^- \bar{\Xi}^+$	$(4.7 \pm 0.7) \times 10^{-4}$	1081		
$K^+ K^- \pi^+ \pi^- \pi^0$	$(8.6 \pm 0.9) \times 10^{-3}$		1545	$\eta_c \pi^+ \pi^-$	$< 7 \times 10^{-4}$	CL=90% 308		
$K_S^0 K^\pm \pi^\mp \pi^+ \pi^-$	$(4.2 \pm 0.4) \times 10^{-3}$		1544	<b>Radiative decays</b>				
$K^+ K^- \pi^0 \pi^0$	$(5.4 \pm 0.9) \times 10^{-3}$		1582	$\gamma J/\psi(1S)$	$(1.27 \pm 0.06) \%$	303		
$K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.}$	$(2.44 \pm 0.33) \%$		1581	$\gamma \rho^0$	$< 9 \times 10^{-6}$	CL=90% 1619		
$\rho^+ K^- K^0 + \text{c.c.}$	$(1.18 \pm 0.21) \%$		1458	$\gamma \omega$	$< 8 \times 10^{-6}$	CL=90% 1618		
$K^*(892)^- K^+ \pi^0 \rightarrow$	$(4.5 \pm 1.1) \times 10^{-3}$	-		$\gamma \phi$	$< 6 \times 10^{-6}$	CL=90% 1555		
$K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.}$				$\gamma \gamma$	$(2.23 \pm 0.13) \times 10^{-4}$	1707		
<div><div><math>\chi_{c1}(1P)</math></div><div><math>I^G(J^{PC}) = 0^+(1^{++})</math></div><div>Mass <math>m = 3510.66 \pm 0.07</math> MeV (S = 1.5)</div><div>Full width <math>\Gamma = 0.84 \pm 0.04</math> MeV</div></div>								

# Meson Summary Table

$\chi_{c1}(1P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
Hadronic decays			
$3(\pi^+\pi^-)$	$(5.8 \pm 1.4) \times 10^{-3}$	S=1.2	1683
$2(\pi^+\pi^-)$	$(7.6 \pm 2.6) \times 10^{-3}$		1728
$\pi^+\pi^-\pi^0\pi^0$	$(1.22 \pm 0.16) \%$		1729
$\rho^+\pi^-\pi^0 + \text{c.c.}$	$(1.48 \pm 0.25) \%$		1658
$\rho^0\pi^+\pi^-$	$(3.9 \pm 3.5) \times 10^{-3}$		1657
$4\pi^0$	$(5.5 \pm 0.8) \times 10^{-4}$		1729
$\pi^+\pi^-K^+K^-$	$(4.5 \pm 1.0) \times 10^{-3}$		1632
$K^+K^-\pi^0\pi^0$	$(1.14 \pm 0.28) \times 10^{-3}$		1634
$K^+K^-\pi^+\pi^-\pi^0$	$(1.15 \pm 0.13) \%$		1598
$K_S^0K^\pm\pi^\mp\pi^+\pi^-$	$(7.5 \pm 0.8) \times 10^{-3}$		1596
$K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	$(8.7 \pm 1.4) \times 10^{-3}$		1632
$\rho^-K^+\bar{K}^0 + \text{c.c.}$	$(5.1 \pm 1.2) \times 10^{-3}$		1514
$K^*(892)^0\bar{K}^0\pi^0 \rightarrow$ $K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	$(2.4 \pm 0.7) \times 10^{-3}$		—
$K^+K^-\eta\pi^0$	$(1.14 \pm 0.35) \times 10^{-3}$		1523
$\pi^+\pi^-K_S^0K_S^0$	$(7.0 \pm 3.0) \times 10^{-4}$		1630
$K^+K^-\eta$	$(3.2 \pm 1.0) \times 10^{-4}$		1566
$\bar{K}^0K^+\pi^- + \text{c.c.}$	$(7.1 \pm 0.6) \times 10^{-3}$		1661
$K^*(892)^0\bar{K}^0 + \text{c.c.}$	$(1.0 \pm 0.4) \times 10^{-3}$		1602
$K^*(892)^+K^- + \text{c.c.}$	$(1.5 \pm 0.7) \times 10^{-3}$		1602
$K_J^*(1430)^0\bar{K}^0 + \text{c.c.} \rightarrow$ $K_S^0K^+\pi^- + \text{c.c.}$	$< 8 \times 10^{-4}$	CL=90%	—
$K_J^*(1430)^+K^- + \text{c.c.} \rightarrow$ $K_S^0K^+\pi^- + \text{c.c.}$	$< 2.2 \times 10^{-3}$	CL=90%	—
$K^+K^-\pi^0$	$(1.85 \pm 0.25) \times 10^{-3}$		1662
$\eta\pi^+\pi^-$	$(4.9 \pm 0.5) \times 10^{-3}$		1701
$a_0(980)^+\pi^- + \text{c.c.} \rightarrow \eta\pi^+\pi^-$	$(1.8 \pm 0.6) \times 10^{-3}$		—
$f_2(1270)\eta$	$(2.7 \pm 0.8) \times 10^{-3}$		1467
$\pi^+\pi^-\eta'$	$(2.3 \pm 0.5) \times 10^{-3}$		1612
$K^+K^-\eta'(958)$	$(8.8 \pm 0.9) \times 10^{-4}$		1461
$K_0^*(1430)+K^- + \text{c.c.}$	$(6.4 \pm_{-2.8}^{+2.2}) \times 10^{-4}$		—
$f_0(980)\eta'(958)$	$(1.6 \pm_{-0.7}^{+1.4}) \times 10^{-4}$		1460
$f_0(1710)\eta'(958)$	$(7 \pm_{-5}^{+7}) \times 10^{-5}$		1106
$f_2'(1525)\eta'(958)$	$(9 \pm 6) \times 10^{-5}$		1225
$\pi^0f_0(980) \rightarrow \pi^0\pi^+\pi^-$	$< 6 \times 10^{-6}$	CL=90%	—
$K^+\bar{K}^*(892)^0\pi^- + \text{c.c.}$	$(3.2 \pm 2.1) \times 10^{-3}$		1577
$K^*(892)^0\bar{K}^*(892)^0$	$(1.5 \pm 0.4) \times 10^{-3}$		1512
$K^+K^-K_S^0K_S^0$	$< 4 \times 10^{-4}$	CL=90%	1390
$K^+K^-K^+K^-$	$(5.5 \pm 1.1) \times 10^{-4}$		1393
$K^+K^-\phi$	$(4.2 \pm 1.6) \times 10^{-4}$		1440
$\bar{K}^0K^+\pi^-\phi + \text{c.c.}$	$(3.3 \pm 0.5) \times 10^{-3}$		1387
$K^+K^-\pi^0\phi$	$(1.62 \pm 0.30) \times 10^{-3}$		1390
$\phi\pi^+\pi^-\pi^0$	$(7.5 \pm 1.0) \times 10^{-4}$		1578
$\omega\omega$	$(5.8 \pm 0.7) \times 10^{-4}$		1571
$\omega K^+K^-$	$(7.8 \pm 0.9) \times 10^{-4}$		1513
$\omega\phi$	$(2.1 \pm 0.6) \times 10^{-5}$		1503
$\phi\phi$	$(4.2 \pm 0.5) \times 10^{-4}$		1429
$p\bar{p}$	$(7.72 \pm 0.35) \times 10^{-5}$		1484
$p\bar{p}\pi^0$	$(1.59 \pm 0.19) \times 10^{-4}$		1438
$p\bar{p}\eta$	$(1.48 \pm 0.25) \times 10^{-4}$		1254
$p\bar{p}\omega$	$(2.16 \pm 0.31) \times 10^{-4}$		1117
$p\bar{p}\phi$	$< 1.8 \times 10^{-5}$	CL=90%	962
$p\bar{p}\pi^+\pi^-$	$(5.0 \pm 1.9) \times 10^{-4}$		1381
$p\bar{p}K^+K^-$ (non-resonant)	$(1.30 \pm 0.23) \times 10^{-4}$		974
$p\bar{p}K_S^0K_S^0$	$< 4.5 \times 10^{-4}$	CL=90%	968
$p\bar{n}\pi^-$	$(3.9 \pm 0.5) \times 10^{-4}$		1435
$\bar{p}n\pi^+$	$(4.0 \pm 0.5) \times 10^{-4}$		1435
$p\bar{n}\pi^-\pi^0$	$(1.05 \pm 0.12) \times 10^{-3}$		1383
$\bar{p}n\pi^+\pi^0$	$(1.03 \pm 0.12) \times 10^{-3}$		1383
$\Lambda\bar{\Lambda}$	$(1.16 \pm 0.12) \times 10^{-4}$		1355
$\Lambda\bar{\Lambda}\pi^+\pi^-$	$(3.0 \pm 0.5) \times 10^{-4}$		1223
$\Lambda\bar{\Lambda}\pi^+\pi^-$ (non-resonant)	$(2.5 \pm 0.6) \times 10^{-4}$		1223
$\Sigma(1385)^+\bar{\Lambda}\pi^- + \text{c.c.}$	$< 1.3 \times 10^{-4}$	CL=90%	1157
$\Sigma(1385)^-\bar{\Lambda}\pi^+ + \text{c.c.}$	$< 1.3 \times 10^{-4}$	CL=90%	1157
$K^+\bar{p}\Lambda$	$(4.2 \pm 0.4) \times 10^{-4}$	S=1.1	1203
$K^+\bar{p}\Lambda(1520) + \text{c.c.}$	$(1.7 \pm 0.5) \times 10^{-4}$		950
$\Lambda(1520)\bar{\Lambda}(1520)$	$< 1.0 \times 10^{-4}$	CL=90%	879
$\Sigma^0\bar{\Sigma}^0$	$< 4 \times 10^{-5}$	CL=90%	1288
$\Sigma^+\bar{\Sigma}^-$	$< 6 \times 10^{-5}$	CL=90%	1291
$\Sigma(1385)^+\bar{\Sigma}(1385)^-$	$< 1.0 \times 10^{-4}$	CL=90%	1081
$\Sigma(1385)^-\bar{\Sigma}(1385)^+$	$< 5 \times 10^{-5}$	CL=90%	1081

$K^-\Lambda\bar{\Xi}^++ \text{ c.c.}$	$(1.38 \pm 0.25) \times 10^{-4}$		963
$\Xi^0\Xi^0$	$< 6 \times 10^{-5}$	CL=90%	1163
$\Xi^-\Xi^++$	$(8.2 \pm 2.2) \times 10^{-5}$		1155
$\pi^+\pi^- + K^+K^-$	$< 2.1 \times 10^{-3}$		—
$K_S^0K_S^0$	$< 6 \times 10^{-5}$	CL=90%	1683
$\eta_c\pi^+\pi^-$	$< 3.2 \times 10^{-3}$	CL=90%	413

## Radiative decays

$\gamma J/\psi(1S)$	$(33.9 \pm 1.2) \%$		389
$\gamma\rho^0$	$(2.20 \pm 0.18) \times 10^{-4}$		1670
$\gamma\omega$	$(6.9 \pm 0.8) \times 10^{-5}$		1668
$\gamma\phi$	$(2.5 \pm 0.5) \times 10^{-5}$		1607

$h_c(1P)$

$I^G(J^{PC}) = ?^?(1^{+-})$

Mass  $m = 3525.38 \pm 0.11$  MeV

Full width  $\Gamma = 0.7 \pm 0.4$  MeV

$h_c(1P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$J/\psi(1S)\pi\pi$	not seen		312
$p\bar{p}$	$< 1.5 \times 10^{-4}$	90%	1492
$\eta_c(1S)\gamma$	$(51 \pm 6) \%$		500
$\pi^+\pi^-\pi^0$	$< 2.2 \times 10^{-3}$		1749
$2\pi^+2\pi^-\pi^0$	$(2.2 \pm_{-0.7}^{+0.8}) \%$		1716
$3\pi^+3\pi^-\pi^0$	$< 2.9 \%$		1661

$\chi_{c2}(1P)$

$I^G(J^{PC}) = 0^+(2^{++})$

Mass  $m = 3556.20 \pm 0.09$  MeV

Full width  $\Gamma = 1.93 \pm 0.11$  MeV

$\chi_{c2}(1P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
Hadronic decays			
$2(\pi^+\pi^-)$	$(1.07 \pm 0.10) \%$		1751
$\pi^+\pi^-\pi^0\pi^0$	$(1.91 \pm 0.25) \%$		1752
$\rho^+\pi^-\pi^0 + \text{c.c.}$	$(2.3 \pm 0.4) \%$		1682
$4\pi^0$	$(1.16 \pm 0.16) \times 10^{-3}$		1752
$K^+K^-\pi^0\pi^0$	$(2.2 \pm 0.4) \times 10^{-3}$		1658
$K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	$(1.44 \pm 0.21) \%$		1657
$\rho^-K^+\bar{K}^0 + \text{c.c.}$	$(4.3 \pm 1.3) \times 10^{-3}$		1540
$K^*(892)^0K^-\pi^+ \rightarrow$ $K^-\pi^+K^0\pi^0 + \text{c.c.}$	$(3.1 \pm 0.8) \times 10^{-3}$		—
$K^*(892)^0\bar{K}^0\pi^0 \rightarrow$ $K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	$(4.0 \pm 0.9) \times 10^{-3}$		—
$K^*(892)^-K^+\pi^0 \rightarrow$ $K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	$(3.9 \pm 0.9) \times 10^{-3}$		—
$K^*(892)^+\bar{K}^0\pi^- \rightarrow$ $K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	$(3.1 \pm 0.8) \times 10^{-3}$		—
$K^+K^-\eta\pi^0$	$(1.3 \pm 0.5) \times 10^{-3}$		1549
$K^+K^-\pi^+\pi^-$	$(8.9 \pm 1.0) \times 10^{-3}$		1656
$K^+K^-\pi^+\pi^-\pi^0$	$(1.17 \pm 0.13) \%$		1623
$K_S^0K^\pm\pi^\mp\pi^+\pi^-$	$(7.3 \pm 0.8) \times 10^{-3}$		1621
$K^+\bar{K}^*(892)^0\pi^- + \text{c.c.}$	$(2.2 \pm 1.1) \times 10^{-3}$		1602
$K^*(892)^0\bar{K}^*(892)^0$	$(2.4 \pm 0.5) \times 10^{-3}$		1538
$3(\pi^+\pi^-)$	$(8.6 \pm 1.8) \times 10^{-3}$		1707
$\phi\phi$	$(1.12 \pm 0.10) \times 10^{-3}$		1457
$\omega\omega$	$(8.8 \pm 1.1) \times 10^{-4}$		1597
$\omega K^+K^-$	$(7.3 \pm 0.9) \times 10^{-4}$		1540
$\pi\pi$	$(2.33 \pm 0.12) \times 10^{-3}$		1773
$\rho^0\pi^+\pi^-$	$(3.8 \pm 1.6) \times 10^{-3}$		1682
$\pi^+\pi^-\eta$	$(5.0 \pm 1.3) \times 10^{-4}$		1724
$\pi^+\pi^-\eta'$	$(5.2 \pm 1.9) \times 10^{-4}$		1636
$\eta\eta$	$(5.7 \pm 0.5) \times 10^{-4}$		1692
$K^+K^-$	$(1.05 \pm 0.07) \times 10^{-3}$		1708
$K_S^0K_S^0$	$(5.5 \pm 0.4) \times 10^{-4}$		1707
$\bar{K}^0K^+\pi^- + \text{c.c.}$	$(1.34 \pm 0.19) \times 10^{-3}$		1685
$K^+K^-\pi^0$	$(3.2 \pm 0.8) \times 10^{-4}$		1686
$K^+K^-\eta$	$< 3.4 \times 10^{-4}$	90%	1592
$K^+K^-\eta'(958)$	$(1.94 \pm 0.34) \times 10^{-4}$		1488
$\eta\eta'$	$< 6 \times 10^{-5}$	90%	1600
$\eta'\eta'$	$< 1.0 \times 10^{-4}$	90%	1498
$\pi^+\pi^-K_S^0K_S^0$	$(2.3 \pm 0.6) \times 10^{-3}$		1655
$K^+K^-K_S^0K_S^0$	$< 4 \times 10^{-4}$	90%	1418
$K^+K^-K^+K^-$	$(1.73 \pm 0.21) \times 10^{-3}$		1421

## Meson Summary Table

$K^+ K^- \phi$	$(1.48 \pm 0.31) \times 10^{-3}$		1468
$\bar{K}^0 K^+ \pi^- \phi + \text{c.c.}$	$(4.8 \pm 0.7) \times 10^{-3}$		1416
$K^+ K^- \pi^0 \phi$	$(2.7 \pm 0.5) \times 10^{-3}$		1419
$\phi \pi^+ \pi^- \pi^0$	$(9.3 \pm 1.2) \times 10^{-4}$		1603
$\rho \bar{\rho}$	$(7.5 \pm 0.4) \times 10^{-5}$		1510
$\rho \bar{\rho} \pi^0$	$(4.9 \pm 0.4) \times 10^{-4}$		1465
$\rho \bar{\rho} \eta$	$(1.82 \pm 0.26) \times 10^{-4}$		1285
$\rho \bar{\rho} \omega$	$(3.8 \pm 0.5) \times 10^{-4}$		1152
$\rho \bar{\rho} \phi$	$(2.9 \pm 0.9) \times 10^{-5}$		1002
$\rho \bar{\rho} \pi^+ \pi^-$	$(1.32 \pm 0.34) \times 10^{-3}$		1410
$\rho \bar{\rho} \pi^0 \pi^0$	$(8.2 \pm 2.5) \times 10^{-4}$		1414
$\rho \bar{\rho} K^+ K^- (\text{non-resonant})$	$(2.00 \pm 0.34) \times 10^{-4}$		1013
$\rho \bar{\rho} K_S^0 K_S^0$	$< 7.9 \times 10^{-4}$	90%	1007
$\rho \bar{\rho} \pi^-$	$(8.9 \pm 1.0) \times 10^{-4}$		1463
$\bar{\rho} n \pi^+$	$(9.3 \pm 0.9) \times 10^{-4}$		1463
$\rho \bar{\rho} \pi^- \pi^0$	$(2.27 \pm 0.19) \times 10^{-3}$		1411
$\bar{\rho} n \pi^+ \pi^0$	$(2.21 \pm 0.20) \times 10^{-3}$		1411
$\Lambda \bar{\Lambda}$	$(1.92 \pm 0.16) \times 10^{-4}$		1385
$\Lambda \bar{\Lambda} \pi^+ \pi^-$	$(1.31 \pm 0.17) \times 10^{-3}$		1255
$\Lambda \bar{\Lambda} \pi^+ \pi^- (\text{non-resonant})$	$(6.9 \pm 1.6) \times 10^{-4}$		1255
$\Sigma(1385)^+ \bar{\Lambda} \pi^- + \text{c.c.}$	$< 4 \times 10^{-4}$	90%	1192
$\Sigma(1385)^- \bar{\Lambda} \pi^+ + \text{c.c.}$	$< 6 \times 10^{-4}$	90%	1192
$K^+ \bar{\rho} \Lambda + \text{c.c.}$	$(8.1 \pm 0.6) \times 10^{-4}$		1236
$K^+ \bar{\rho} \Lambda(1520) + \text{c.c.}$	$(2.9 \pm 0.7) \times 10^{-4}$		992
$\Lambda(1520) \bar{\Lambda}(1520)$	$(4.8 \pm 1.5) \times 10^{-4}$		923
$\Sigma^0 \bar{\Sigma}^0$	$< 6 \times 10^{-5}$	90%	1319
$\Sigma^+ \bar{\Sigma}^-$	$< 7 \times 10^{-5}$	90%	1322
$\Sigma(1385)^+ \bar{\Sigma}(1385)^-$	$< 1.6 \times 10^{-4}$	90%	1118
$\Sigma(1385)^- \bar{\Sigma}(1385)^+$	$< 8 \times 10^{-5}$	90%	1118
$K^- \Lambda \bar{\Xi}^+ + \text{c.c.}$	$(1.84 \pm 0.34) \times 10^{-4}$		1004
$\Xi^0 \bar{\Xi}^0$	$< 1.1 \times 10^{-4}$	90%	1197
$\Xi^- \bar{\Xi}^+$	$(1.48 \pm 0.33) \times 10^{-4}$		1189
$J/\psi(1S) \pi^+ \pi^- \pi^0$	$< 1.5 \%$	90%	185
$\pi^0 \eta_c$	$< 3.2 \times 10^{-3}$	90%	512
$\eta_c(1S) \pi^+ \pi^-$	$< 5.4 \times 10^{-3}$	90%	459

## Radiative decays

$\gamma J/\psi(1S)$	$(19.2 \pm 0.7) \%$		430
$\gamma \rho^0$	$< 2.0 \times 10^{-5}$	90%	1694
$\gamma \omega$	$< 6 \times 10^{-6}$	90%	1692
$\gamma \phi$	$< 8 \times 10^{-6}$	90%	1632
$\gamma \gamma$	$(2.74 \pm 0.14) \times 10^{-4}$		1778

 **$\eta_c(2S)$** 

$$J^{PC} = 0^+(0^-+)$$

Quantum numbers are quark model predictions.

Mass  $m = 3639.2 \pm 1.2$  MeV

Full width  $\Gamma = 11.3^{+3.2}_{-2.9}$  MeV

$\eta_c(2S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
hadrons	not seen		—
$K \bar{K} \pi$	$(1.9 \pm 1.2) \%$		1730
$K \bar{K} \eta$	$(5 \pm 4) \times 10^{-3}$		1638
$2\pi^+ 2\pi^-$	not seen		1793
$\rho^0 \rho^0$	not seen		1646
$3\pi^+ 3\pi^-$	not seen		1750
$K^+ K^- \pi^+ \pi^-$	not seen		1701
$K^{*0} \bar{K}^{*0}$	not seen		1586
$K^+ K^- \pi^+ \pi^- \pi^0$	$(1.4 \pm 1.0) \%$		1668
$K^+ K^- 2\pi^+ 2\pi^-$	not seen		1628
$K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.}$	seen		1667
$2K^+ 2K^-$	not seen		1471
$\phi \phi$	not seen		1507
$\rho \bar{\rho}$	$< 2.0 \times 10^{-3}$	90%	1559
$\gamma \gamma$	$(1.9 \pm 1.3) \times 10^{-4}$		1820
$\pi^+ \pi^- \eta$	not seen		1767
$\pi^+ \pi^- \eta'$	not seen		1681
$\pi^+ \pi^- \eta_c(1S)$	$< 25 \%$	90%	539

 **$\psi(2S)$** 

$$J^{PC} = 0^-(1^{--})$$

Mass  $m = 3686.097 \pm 0.025$  MeV ( $S = 2.6$ )

Full width  $\Gamma = 296 \pm 8$  keV

$\Gamma_{ee} = 2.34 \pm 0.04$  keV

$\psi(2S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
hadrons	$(97.85 \pm 0.13) \%$		—
virtual $\gamma \rightarrow$ hadrons	$(1.73 \pm 0.14) \%$	$S=1.5$	—
$g g g$	$(10.6 \pm 1.6) \%$		—
$\gamma g g$	$(1.03 \pm 0.29) \%$		—
light hadrons	$(15.4 \pm 1.5) \%$		—
$e^+ e^-$	$(7.89 \pm 0.17) \times 10^{-3}$		1843
$\mu^+ \mu^-$	$(7.9 \pm 0.9) \times 10^{-3}$		1840
$\tau^+ \tau^-$	$(3.1 \pm 0.4) \times 10^{-3}$		489

Decays into  $J/\psi(1S)$  and anything

$J/\psi(1S)$ anything	$(61.0 \pm 0.6) \%$		—
$J/\psi(1S)$ neutrals	$(25.14 \pm 0.33) \%$		—
$J/\psi(1S) \pi^+ \pi^-$	$(34.49 \pm 0.30) \%$		477
$J/\psi(1S) \pi^0 \pi^0$	$(18.16 \pm 0.31) \%$		481
$J/\psi(1S) \eta$	$(3.36 \pm 0.05) \%$		199
$J/\psi(1S) \pi^0$	$(1.268 \pm 0.032) \times 10^{-3}$		528

## Hadronic decays

$\pi^0 h_c(1P)$	$(8.6 \pm 1.3) \times 10^{-4}$		85
$3(\pi^+ \pi^-) \pi^0$	$(3.5 \pm 1.6) \times 10^{-3}$		1746
$2(\pi^+ \pi^-) \pi^0$	$(2.9 \pm 1.0) \times 10^{-3}$	$S=4.7$	1799
$\rho a_2(1320)$	$(2.6 \pm 0.9) \times 10^{-4}$		1500
$\rho \bar{\rho}$	$(2.88 \pm 0.09) \times 10^{-4}$		1586
$\Delta^{++} \bar{\Delta}^{--}$	$(1.28 \pm 0.35) \times 10^{-4}$		1371
$\Lambda \bar{\Lambda} \pi^0$	$< 2.9 \times 10^{-6}$	CL=90%	1412
$\Lambda \bar{\Lambda} \eta$	$(2.5 \pm 0.4) \times 10^{-5}$		1197
$\Lambda \bar{\rho} K^+$	$(1.00 \pm 0.14) \times 10^{-4}$		1327
$\Lambda \bar{\rho} K^+ \pi^+ \pi^-$	$(1.8 \pm 0.4) \times 10^{-4}$		1167
$\Lambda \bar{\Lambda} \pi^+ \pi^-$	$(2.8 \pm 0.6) \times 10^{-4}$		1346
$\Lambda \bar{\Lambda}$	$(3.57 \pm 0.18) \times 10^{-4}$		1467
$\Lambda \bar{\Sigma}^+ \pi^- + \text{c.c.}$	$(1.40 \pm 0.13) \times 10^{-4}$		1376
$\Lambda \bar{\Sigma}^- \pi^+ + \text{c.c.}$	$(1.54 \pm 0.14) \times 10^{-4}$		1379
$\Sigma^0 \bar{\rho} K^+ + \text{c.c.}$	$(1.67 \pm 0.18) \times 10^{-5}$		1291
$\Sigma^+ \bar{\Sigma}^-$	$(2.51 \pm 0.21) \times 10^{-4}$		1408
$\Sigma^0 \bar{\Sigma}^0$	$(2.32 \pm 0.16) \times 10^{-4}$		1405
$\Sigma(1385)^+ \bar{\Sigma}(1385)^-$	$(1.1 \pm 0.4) \times 10^{-4}$		1218
$\Xi^- \bar{\Xi}^+$	$(2.64 \pm 0.18) \times 10^{-4}$		1284
$\Xi^0 \bar{\Xi}^0$	$(2.07 \pm 0.23) \times 10^{-4}$		1291
$\Xi(1530)^0 \bar{\Xi}(1530)^0$	$(5.2 \pm 3.2) \times 10^{-5}$		1025
$K^- \Lambda \bar{\Xi}^+ + \text{c.c.}$	$(3.9 \pm 0.4) \times 10^{-5}$		1114
$\Xi(1690)^- \bar{\Xi}^+ \rightarrow K^- \Lambda \bar{\Xi}^+ + \text{c.c.}$	$(5.2 \pm 1.6) \times 10^{-6}$		—
$\Xi(1820)^- \bar{\Xi}^+ \rightarrow K^- \Lambda \bar{\Xi}^+ + \text{c.c.}$	$(1.20 \pm 0.32) \times 10^{-5}$		—
$K^- \Sigma^0 \bar{\Xi}^+ + \text{c.c.}$	$(3.7 \pm 0.4) \times 10^{-5}$		1060
$\Omega^- \bar{\Omega}^+$	$(4.7 \pm 1.0) \times 10^{-5}$		774
$\pi^0 \rho \bar{\rho}$	$(1.53 \pm 0.07) \times 10^{-4}$		1543
$N(940) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(6.4 \pm 1.8) \times 10^{-5}$		—
$N(1440) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(7.3 \pm 1.7) \times 10^{-5}$	$S=2.5$	—
$N(1520) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(6.4 \pm 2.3) \times 10^{-6}$		—
$N(1535) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(2.5 \pm 1.0) \times 10^{-5}$		—
$N(1650) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(3.8 \pm 1.4) \times 10^{-5}$		—
$N(1720) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(1.79 \pm 0.26) \times 10^{-5}$		—
$N(2300) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(2.6 \pm 1.2) \times 10^{-5}$		—
$N(2570) \bar{\rho} + \text{c.c.} \rightarrow \pi^0 \rho \bar{\rho}$	$(2.13 \pm 0.40) \times 10^{-5}$		—
$\pi^0 f_0(2100) \rightarrow \pi^0 \rho \bar{\rho}$	$(1.1 \pm 0.4) \times 10^{-5}$		—
$\eta \rho \bar{\rho}$	$(6.0 \pm 0.4) \times 10^{-5}$		1373
$\eta f_0(2100) \rightarrow \eta \rho \bar{\rho}$	$(1.2 \pm 0.4) \times 10^{-5}$		—
$N(1535) \bar{\rho} \rightarrow \eta \rho \bar{\rho}$	$(4.4 \pm 0.7) \times 10^{-5}$		—
$\omega \rho \bar{\rho}$	$(6.9 \pm 2.1) \times 10^{-5}$		1247
$\phi \rho \bar{\rho}$	$< 2.4 \times 10^{-5}$	CL=90%	1109
$\pi^+ \pi^- \rho \bar{\rho}$	$(6.0 \pm 0.4) \times 10^{-4}$		1491
$\rho \bar{\rho} \pi^- \text{ or c.c.}$	$(2.48 \pm 0.17) \times 10^{-4}$		—
$\rho \bar{\rho} \pi^- \pi^0$	$(3.2 \pm 0.7) \times 10^{-4}$		1492
$2(\pi^+ \pi^- \pi^0)$	$(4.8 \pm 1.5) \times 10^{-3}$		1776
$\eta \pi^+ \pi^-$	$< 1.6 \times 10^{-4}$	CL=90%	1791
$\eta \pi^+ \pi^- \pi^0$	$(9.5 \pm 1.7) \times 10^{-4}$		1778
$2(\pi^+ \pi^-) \eta$	$(1.2 \pm 0.6) \times 10^{-3}$		1758
$\eta' \pi^+ \pi^- \pi^0$	$(4.5 \pm 2.1) \times 10^{-4}$		1692
$\omega \pi^+ \pi^-$	$(7.3 \pm 1.2) \times 10^{-4}$	$S=2.1$	1748
$b_1^\pm \pi^\mp$	$(4.0 \pm 0.6) \times 10^{-4}$	$S=1.1$	1635

$b_1^0 \pi^0$	( 2.4 ± 0.6 ) × 10 <sup>-4</sup>	—	$\gamma \eta_c(2S)$	( 7 ± 5 ) × 10 <sup>-4</sup>	47
$\omega f_2(1270)$	( 2.2 ± 0.4 ) × 10 <sup>-4</sup>	1515	$\gamma \pi_0^0$	( 1.6 ± 0.4 ) × 10 <sup>-6</sup>	1841
$\pi^+ \pi^- K^+ K^-$	( 7.5 ± 0.9 ) × 10 <sup>-4</sup>	S=1.9 1726	$\gamma \eta'(958)$	( 1.23 ± 0.06 ) × 10 <sup>-4</sup>	1719
$\rho^0 K^+ K^-$	( 2.2 ± 0.4 ) × 10 <sup>-4</sup>	1616	$\gamma f_2(1270)$	( 2.73 $^{+0.29}_{-0.25}$ ) × 10 <sup>-4</sup>	S=1.8 1622
$K^*(892)^0 \bar{K}_2^*(1430)^0$	( 1.9 ± 0.5 ) × 10 <sup>-4</sup>	1418	$\gamma f_0(1370) \rightarrow \gamma K \bar{K}$	( 3.1 ± 1.7 ) × 10 <sup>-5</sup>	1588
$K^+ K^- \pi^+ \pi^- \eta$	( 1.3 ± 0.7 ) × 10 <sup>-3</sup>	1574	$\gamma f_0(1500)$	( 9.2 ± 1.9 ) × 10 <sup>-5</sup>	1536
$K^+ K^- 2(\pi^+ \pi^-) \pi^0$	( 1.00 ± 0.31 ) × 10 <sup>-3</sup>	1611	$\gamma f_2'(1525)$	( 3.3 ± 0.8 ) × 10 <sup>-5</sup>	1528
$K^+ K^- 2(\pi^+ \pi^-)$	( 1.9 ± 0.9 ) × 10 <sup>-3</sup>	1654	$\gamma f_0(1710) \rightarrow \gamma \pi \pi$	( 3.5 ± 0.6 ) × 10 <sup>-5</sup>	—
$K_1(1270)^\pm K^\mp$	( 1.00 ± 0.28 ) × 10 <sup>-3</sup>	1581	$\gamma f_0(1710) \rightarrow \gamma K \bar{K}$	( 6.6 ± 0.7 ) × 10 <sup>-5</sup>	—
$K_S^0 K_S^0 \pi^+ \pi^-$	( 2.2 ± 0.4 ) × 10 <sup>-4</sup>	1724	$\gamma f_0(2100) \rightarrow \gamma \pi \pi$	( 4.8 ± 1.0 ) × 10 <sup>-6</sup>	1244
$\rho^0 p \bar{p}$	( 5.0 ± 2.2 ) × 10 <sup>-5</sup>	1252	$\gamma f_0(2200) \rightarrow \gamma K \bar{K}$	( 3.2 ± 1.0 ) × 10 <sup>-6</sup>	1193
$K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.}$	( 6.7 ± 2.5 ) × 10 <sup>-4</sup>	1674	$\gamma f_0(2220) \rightarrow \gamma \pi \pi$	< 5.8 × 10 <sup>-6</sup>	CL=90% 1168
$2(\pi^+ \pi^-)$	( 2.4 ± 0.6 ) × 10 <sup>-4</sup>	S=2.2 1817	$\gamma f_J(2220) \rightarrow \gamma K \bar{K}$	< 9.5 × 10 <sup>-6</sup>	CL=90% 1168
$\rho^0 \pi^+ \pi^-$	( 2.2 ± 0.6 ) × 10 <sup>-4</sup>	S=1.4 1750	$\gamma \gamma$	< 1.5 × 10 <sup>-4</sup>	CL=90% 1843
$K^+ K^- \pi^+ \pi^- \pi^0$	( 1.26 ± 0.09 ) × 10 <sup>-3</sup>	1694	$\gamma \eta$	( 1.4 ± 0.5 ) × 10 <sup>-6</sup>	1802
$\omega f_0(1710) \rightarrow \omega K^+ K^-$	( 5.9 ± 2.2 ) × 10 <sup>-5</sup>	—	$\gamma \eta \pi^+ \pi^-$	( 8.7 ± 2.1 ) × 10 <sup>-4</sup>	1791
$K^*(892)^0 K^- \pi^+ \pi^0 + \text{c.c.}$	( 8.6 ± 2.2 ) × 10 <sup>-4</sup>	—	$\gamma \eta(1405) \rightarrow \gamma K \bar{K} \pi$	< 9 × 10 <sup>-5</sup>	CL=90% 1569
$K^*(892)^+ K^- \pi^+ \pi^- + \text{c.c.}$	( 9.6 ± 2.8 ) × 10 <sup>-4</sup>	—	$\gamma \eta(1405) \rightarrow \eta \pi^+ \pi^-$	( 3.6 ± 2.5 ) × 10 <sup>-5</sup>	—
$K^*(892)^+ K^- \rho^0 + \text{c.c.}$	( 7.3 ± 2.6 ) × 10 <sup>-4</sup>	—	$\gamma \eta(1475) \rightarrow K \bar{K} \pi$	< 1.4 × 10 <sup>-4</sup>	CL=90% —
$K^*(892)^0 K^- \rho^+ + \text{c.c.}$	( 6.1 ± 1.8 ) × 10 <sup>-4</sup>	—	$\gamma \eta(1475) \rightarrow \eta \pi^+ \pi^-$	< 8.8 × 10 <sup>-5</sup>	CL=90% —
$\eta K^+ K^-$ , no $\eta \phi$	( 3.1 ± 0.4 ) × 10 <sup>-5</sup>	1664	$\gamma 2(\pi^+ \pi^-)$	( 4.0 ± 0.6 ) × 10 <sup>-4</sup>	1817
$\omega K^+ K^-$	( 1.62 ± 0.11 ) × 10 <sup>-4</sup>	S=1.1 1614	$\gamma K^{*0} K^+ \pi^- + \text{c.c.}$	( 3.7 ± 0.9 ) × 10 <sup>-4</sup>	1674
$\omega K^*(892)^+ K^- + \text{c.c.}$	( 2.07 ± 0.26 ) × 10 <sup>-4</sup>	1482	$\gamma K^{*0} \bar{K}^{*0}$	( 2.4 ± 0.7 ) × 10 <sup>-4</sup>	1613
$\omega K_2^*(1430)^+ K^- + \text{c.c.}$	( 6.1 ± 1.2 ) × 10 <sup>-5</sup>	1253	$\gamma K_S^0 K^+ \pi^- + \text{c.c.}$	( 2.6 ± 0.5 ) × 10 <sup>-4</sup>	1753
$\omega \bar{K}^*(892)^0 K^0$	( 1.68 ± 0.30 ) × 10 <sup>-4</sup>	1481	$\gamma K^+ K^- \pi^+ \pi^-$	( 1.9 ± 0.5 ) × 10 <sup>-4</sup>	1726
$\omega \bar{K}_2^*(1430)^0 K^0$	( 5.8 ± 2.2 ) × 10 <sup>-5</sup>	1251	$\gamma p \bar{p}$	( 3.9 ± 0.5 ) × 10 <sup>-5</sup>	S=2.0 1586
$\omega X(1440) \rightarrow \omega K_S^0 K^- \pi^+ + \text{c.c.}$	( 1.6 ± 0.4 ) × 10 <sup>-5</sup>	—	$\gamma f_2(1950) \rightarrow \gamma p \bar{p}$	( 1.20 ± 0.22 ) × 10 <sup>-5</sup>	—
$\omega X(1440) \rightarrow \omega K^+ K^- \pi^0$	( 1.09 ± 0.26 ) × 10 <sup>-5</sup>	—	$\gamma f_2(2150) \rightarrow \gamma p \bar{p}$	( 7.2 ± 1.8 ) × 10 <sup>-6</sup>	—
$\omega f_1(1285) \rightarrow \omega K_S^0 K^- \pi^+ + \text{c.c.}$	( 3.0 ± 1.0 ) × 10 <sup>-6</sup>	—	$\gamma X(1835) \rightarrow \gamma p \bar{p}$	( 4.6 $^{+1.8}_{-4.0}$ ) × 10 <sup>-6</sup>	—
$\omega f_1(1285) \rightarrow \omega K^+ K^- \pi^0$	( 1.2 ± 0.7 ) × 10 <sup>-6</sup>	—	$\gamma X \rightarrow \gamma p \bar{p}$	[xxaa] < 2 × 10 <sup>-6</sup>	CL=90% —
$3(\pi^+ \pi^-)$	( 3.5 ± 2.0 ) × 10 <sup>-4</sup>	S=2.8 1774	$\gamma \pi^+ \pi^- p \bar{p}$	( 2.8 ± 1.4 ) × 10 <sup>-5</sup>	1491
$p \bar{p} \pi^+ \pi^- \pi^0$	( 7.3 ± 0.7 ) × 10 <sup>-4</sup>	1435	$\gamma 2(\pi^+ \pi^-) K^+ K^-$	< 2.2 × 10 <sup>-4</sup>	CL=90% 1654
$K^+ K^-$	( 7.5 ± 0.5 ) × 10 <sup>-5</sup>	1776	$\gamma 3(\pi^+ \pi^-)$	< 1.7 × 10 <sup>-4</sup>	CL=90% 1774
$K_S^0 K_L^0$	( 5.34 ± 0.33 ) × 10 <sup>-5</sup>	1775	$\gamma K^+ K^- K^+ K^-$	< 4 × 10 <sup>-5</sup>	CL=90% 1499
$\pi^+ \pi^- \pi^0$	( 2.01 ± 0.17 ) × 10 <sup>-4</sup>	S=1.7 1830	$\gamma \gamma J/\psi$	( 3.1 $^{+1.0}_{-1.2}$ ) × 10 <sup>-4</sup>	542
$\rho(2150) \pi \rightarrow \pi^+ \pi^- \pi^0$	( 1.9 $^{+1.2}_{-0.4}$ ) × 10 <sup>-4</sup>	—			
$\gamma(770) \pi \rightarrow \pi^+ \pi^- \pi^0$	( 3.2 ± 1.2 ) × 10 <sup>-5</sup>	S=1.8 —			
$\pi^+ \pi^-$	( 7.8 ± 2.6 ) × 10 <sup>-6</sup>	1838			
$K_1(1400)^\pm K^\mp$	< 3.1 × 10 <sup>-4</sup>	CL=90% 1532			
$K_2^*(1430)^\pm K^\mp$	( 7.1 $^{+1.3}_{-0.9}$ ) × 10 <sup>-5</sup>	—			
$K^+ K^- \pi^0$	( 4.07 ± 0.31 ) × 10 <sup>-5</sup>	1754			
$K^+ K^*(892)^- + \text{c.c.}$	( 2.9 ± 0.4 ) × 10 <sup>-5</sup>	S=1.2 1698			
$K^*(892)^0 \bar{K}^0 + \text{c.c.}$	( 1.09 ± 0.20 ) × 10 <sup>-4</sup>	1697			
$\phi \pi^+ \pi^-$	( 1.17 ± 0.29 ) × 10 <sup>-4</sup>	S=1.7 1690			
$\phi f_0(980) \rightarrow \pi^+ \pi^-$	( 6.8 ± 2.5 ) × 10 <sup>-5</sup>	S=1.2 —			
$2(K^+ K^-)$	( 6.0 ± 1.4 ) × 10 <sup>-5</sup>	1499			
$\phi K^+ K^-$	( 7.0 ± 1.6 ) × 10 <sup>-5</sup>	1546			
$2(K^+ K^-) \pi^0$	( 1.10 ± 0.28 ) × 10 <sup>-4</sup>	1440			
$\phi \eta$	( 3.10 ± 0.31 ) × 10 <sup>-5</sup>	1654			
$\phi \eta'$	( 3.1 ± 1.6 ) × 10 <sup>-5</sup>	1555			
$\omega \eta'$	( 3.2 $^{+2.5}_{-2.1}$ ) × 10 <sup>-5</sup>	1623			
$\omega \pi^0$	( 2.1 ± 0.6 ) × 10 <sup>-5</sup>	1757			
$\rho \eta'$	( 1.9 $^{+1.7}_{-1.2}$ ) × 10 <sup>-5</sup>	1625			
$\rho \eta$	( 2.2 ± 0.6 ) × 10 <sup>-5</sup>	S=1.1 1717			
$\omega \eta$	< 1.1 × 10 <sup>-5</sup>	CL=90% 1715			
$\phi \pi^0$	< 4 × 10 <sup>-7</sup>	CL=90% 1699			
$\eta_c \pi^+ \pi^- \pi^0$	< 1.0 × 10 <sup>-3</sup>	CL=90% 513			
$p \bar{p} K^+ K^-$	( 2.7 ± 0.7 ) × 10 <sup>-5</sup>	1118			
$\bar{\Lambda} n K_S^0 + \text{c.c.}$	( 8.1 ± 1.8 ) × 10 <sup>-5</sup>	1324			
$\phi f_2'(1525)$	( 4.4 ± 1.6 ) × 10 <sup>-5</sup>	1321			
$\Theta(1540) \bar{\Theta}(1540) \rightarrow K_S^0 p K^- \bar{n} + \text{c.c.}$	< 8.8 × 10 <sup>-6</sup>	CL=90% —			
$\Theta(1540) K^- \bar{n} \rightarrow K_S^0 p K^- \bar{n}$	< 1.0 × 10 <sup>-5</sup>	CL=90% —			
$\Theta(1540) K_S^0 \bar{p} \rightarrow K_S^0 \bar{p} K^+ n$	< 7.0 × 10 <sup>-6</sup>	CL=90% —			
$\bar{\Theta}(1540) K^+ n \rightarrow K_S^0 \bar{p} K^+ n$	< 2.6 × 10 <sup>-5</sup>	CL=90% —			
$\bar{\Theta}(1540) K_S^0 p \rightarrow K_S^0 p K^- \bar{n}$	< 6.0 × 10 <sup>-6</sup>	CL=90% —			
$K_S^0 K_S^0$	< 4.6 × 10 <sup>-6</sup>	1775			

## Meson Summary Table

$K_S^0 K_L^0$	< 1.2	$\times 10^{-5}$	CL=90%	1820
$2(\pi^+ \pi^-)$	< 1.12	$\times 10^{-3}$	CL=90%	1861
$2(\pi^+ \pi^-) \pi^0$	< 1.06	$\times 10^{-3}$	CL=90%	1843
$2(\pi^+ \pi^- \pi^0)$	< 5.85	%	CL=90%	1821
$\omega \pi^+ \pi^-$	< 6.0	$\times 10^{-4}$	CL=90%	1794
$3(\pi^+ \pi^-)$	< 9.1	$\times 10^{-3}$	CL=90%	1819
$3(\pi^+ \pi^-) \pi^0$	< 1.37	%	CL=90%	1792
$3(\pi^+ \pi^-) 2\pi^0$	< 11.74	%	CL=90%	1760
$\eta \pi^+ \pi^-$	< 1.24	$\times 10^{-3}$	CL=90%	1836
$\pi^+ \pi^- 2\pi^0$	< 8.9	$\times 10^{-3}$	CL=90%	1862
$\rho^0 \pi^+ \pi^-$	< 6.9	$\times 10^{-3}$	CL=90%	1796
$\eta 3\pi$	< 1.34	$\times 10^{-3}$	CL=90%	1824
$\eta 2(\pi^+ \pi^-)$	< 2.43	%	CL=90%	1804
$\eta \rho^0 \pi^+ \pi^-$	< 1.45	%	CL=90%	1708
$\eta' 3\pi$	< 2.44	$\times 10^{-3}$	CL=90%	1740
$K^+ K^- \pi^+ \pi^-$	< 9.0	$\times 10^{-4}$	CL=90%	1772
$\phi \pi^+ \pi^-$	< 4.1	$\times 10^{-4}$	CL=90%	1737
$K^+ K^- 2\pi^0$	< 4.2	$\times 10^{-3}$	CL=90%	1774
$4(\pi^+ \pi^-)$	< 1.67	%	CL=90%	1757
$4(\pi^+ \pi^-) \pi^0$	< 3.06	%	CL=90%	1720
$\phi f_0(980)$	< 4.5	$\times 10^{-4}$	CL=90%	1597
$K^+ K^- \pi^+ \pi^- \pi^0$	< 2.36	$\times 10^{-3}$	CL=90%	1741
$K^+ K^- \rho^0 \pi^0$	< 8	$\times 10^{-4}$	CL=90%	1624
$K^+ K^- \rho^+ \pi^-$	< 1.46	%	CL=90%	1622
$\omega K^+ K^-$	< 3.4	$\times 10^{-4}$	CL=90%	1664
$\phi \pi^+ \pi^- \pi^0$	< 3.8	$\times 10^{-3}$	CL=90%	1722
$K^{*0} K^- \pi^+ \pi^0 + \text{c.c.}$	< 1.62	%	CL=90%	1693
$K^{*+} K^- \pi^+ \pi^- + \text{c.c.}$	< 3.23	%	CL=90%	1692
$K^+ K^- \pi^+ \pi^- 2\pi^0$	< 2.67	%	CL=90%	1705
$K^+ K^- 2(\pi^+ \pi^-)$	< 1.03	%	CL=90%	1702
$K^+ K^- 2(\pi^+ \pi^-) \pi^0$	< 3.60	%	CL=90%	1660
$\eta K^+ K^-$	< 4.1	$\times 10^{-4}$	CL=90%	1712
$\eta K^+ K^- \pi^+ \pi^-$	< 1.24	%	CL=90%	1624
$\rho^0 K^+ K^-$	< 5.0	$\times 10^{-3}$	CL=90%	1665
$2(K^+ K^-)$	< 6.0	$\times 10^{-4}$	CL=90%	1552
$\phi K^+ K^-$	< 7.5	$\times 10^{-4}$	CL=90%	1598
$2(K^+ K^-) \pi^0$	< 2.9	$\times 10^{-4}$	CL=90%	1493
$2(K^+ K^-) \pi^+ \pi^-$	< 3.2	$\times 10^{-3}$	CL=90%	1425
$K_S^0 K^- \pi^+$	< 3.2	$\times 10^{-3}$	CL=90%	1799
$K_S^0 K^- \pi^+ \pi^0$	< 1.33	%	CL=90%	1773
$K_S^0 K^- \rho^+$	< 6.6	$\times 10^{-3}$	CL=90%	1664
$K_S^0 K^- 2\pi^+ \pi^-$	< 8.7	$\times 10^{-3}$	CL=90%	1739
$K_S^0 K^- \pi^+ \rho^0$	< 1.6	%	CL=90%	1621
$K_S^0 K^- \pi^+ \eta$	< 1.3	%	CL=90%	1669
$K_S^0 K^- 2\pi^+ \pi^- \pi^0$	< 4.18	%	CL=90%	1703
$K_S^0 K^- 2\pi^+ \pi^- \eta$	< 4.8	%	CL=90%	1570
$K_S^0 K^- \pi^+ 2(\pi^+ \pi^-)$	< 1.22	%	CL=90%	1658
$K_S^0 K^- \pi^+ 2\pi^0$	< 2.65	%	CL=90%	1742
$K_S^0 K^- K^+ K^- \pi^+$	< 4.9	$\times 10^{-3}$	CL=90%	1490
$K_S^0 K^- K^+ K^- \pi^+ \pi^0$	< 3.0	%	CL=90%	1427
$K_S^0 K^- K^+ K^- \pi^+ \eta$	< 2.2	%	CL=90%	1214
$K^{*0} K^- \pi^+ + \text{c.c.}$	< 9.7	$\times 10^{-3}$	CL=90%	1722
$p \bar{p} \pi^0$	< 4	$\times 10^{-5}$	CL=90%	1595
$p \bar{p} \pi^+ \pi^-$	< 5.8	$\times 10^{-4}$	CL=90%	1544
$\Lambda \bar{\Lambda}$	< 1.2	$\times 10^{-4}$	CL=90%	1521
$p \bar{p} \pi^+ \pi^- \pi^0$	< 1.85	$\times 10^{-3}$	CL=90%	1490
$\omega p \bar{p}$	< 2.9	$\times 10^{-4}$	CL=90%	1309
$\Lambda \bar{\Lambda} \pi^0$	< 7	$\times 10^{-5}$	CL=90%	1468
$p \bar{p} 2(\pi^+ \pi^-)$	< 2.6	$\times 10^{-3}$	CL=90%	1425
$\eta p \bar{p}$	< 5.4	$\times 10^{-4}$	CL=90%	1430
$\eta p \bar{p} \pi^+ \pi^-$	< 3.3	$\times 10^{-3}$	CL=90%	1284
$\rho^0 p \bar{p}$	< 1.7	$\times 10^{-3}$	CL=90%	1313
$p \bar{p} K^+ K^-$	< 3.2	$\times 10^{-4}$	CL=90%	1185
$\eta p \bar{p} K^+ K^-$	< 6.9	$\times 10^{-3}$	CL=90%	736
$\pi^0 p \bar{p} K^+ K^-$	< 1.2	$\times 10^{-3}$	CL=90%	1093
$\phi p \bar{p}$	< 1.3	$\times 10^{-4}$	CL=90%	1178
$\Lambda \bar{\Lambda} \pi^+ \pi^-$	< 2.5	$\times 10^{-4}$	CL=90%	1404
$\Lambda \bar{p} K^+$	< 2.8	$\times 10^{-4}$	CL=90%	1387
$\Lambda \bar{p} K^+ \pi^+ \pi^-$	< 6.3	$\times 10^{-4}$	CL=90%	1234
$\Lambda \bar{\Lambda} \eta$	< 1.9	$\times 10^{-4}$	CL=90%	1262
$\Sigma^+ \bar{\Sigma}^-$	< 1.0	$\times 10^{-4}$	CL=90%	1464
$\Sigma^0 \bar{\Sigma}^0$	< 4	$\times 10^{-5}$	CL=90%	1462
$\Xi^+ \bar{\Xi}^-$	< 1.5	$\times 10^{-4}$	CL=90%	1346
$\Xi^0 \bar{\Xi}^0$	< 1.4	$\times 10^{-4}$	CL=90%	1353

## Radiative decays

$\gamma \chi_{c2}$	< 6.4	$\times 10^{-4}$	CL=90%	211
$\gamma \chi_{c1}$	( 2.48 $\pm$ 0.23 ) $\times 10^{-3}$			253
$\gamma \chi_{c0}$	( 7.0 $\pm$ 0.6 ) $\times 10^{-3}$			341
$\gamma \eta_c$	< 7	$\times 10^{-4}$	CL=90%	707
$\gamma \eta_c(2S)$	< 9	$\times 10^{-4}$	CL=90%	132
$\gamma \eta'$	< 1.8	$\times 10^{-4}$	CL=90%	1765
$\gamma \eta$	< 1.5	$\times 10^{-4}$	CL=90%	1847
$\gamma \pi^0$	< 2	$\times 10^{-4}$	CL=90%	1884

 **$\psi(3823)$** was  $X(3823)$ .

$$J^G(J^{PC}) = ?(2^{--})$$

$J, P$  need confirmation.

Mass  $m = 3822.2 \pm 1.2$  MeVFull width  $\Gamma < 16$  MeV, CL = 90% **$\psi(3823)$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\chi_{c1} \gamma$	seen	299
$\chi_{c2} \gamma$	not seen	257

 **$X(3872)$** 

$$J^G(J^{PC}) = 0^+(1^{++})$$

Mass  $m = 3871.69 \pm 0.17$  MeV $m_{X(3872)} - m_{J/\psi} = 775 \pm 4$  MeV $m_{X(3872)} - m_{\psi(2S)}$ Full width  $\Gamma < 1.2$  MeV, CL = 90% **$X(3872)$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\pi^+ \pi^- J/\psi(1S)$	> 2.6 %	650
$\omega J/\psi(1S)$	> 1.9 %	†
$D^0 \bar{D}^0 \pi^0$	>32 %	117
$\bar{D}^{*0} D^0$	>24 %	3
$\gamma J/\psi$	> 6 $\times 10^{-3}$	697
$\gamma \psi(2S)$	> 3.0 %	181
$\pi^+ \pi^- \eta_c(1S)$	not seen	746
$p \bar{p}$	not seen	1693

 **$X(3900)$** 

$$J^G(J^{PC}) = 1^+(1^{+-})$$

Mass  $m = 3886.6 \pm 2.4$  MeV (S = 1.6)Full width  $\Gamma = 28.1 \pm 2.6$  MeV **$X(3900)$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$J/\psi \pi$	seen	699
$h_c \pi^\pm$	not seen	318
$\eta_c \pi^\pm \pi^\mp$	not seen	759
$(D \bar{D}^*)^\pm$	seen	—
$D^0 D^{*-} + \text{c.c.}$	seen	150
$D^- D^{*0} + \text{c.c.}$	seen	141
$\omega \pi^\pm$	not seen	1862
$J/\psi \eta$	not seen	509
$D^+ D^{*-} + \text{c.c.}$	seen	—
$D^0 \bar{D}^{*0} + \text{c.c.}$	seen	—

 **$X(3915)$** was  $\chi_{c0}(3915)$ 

$$J^G(J^{PC}) = 0^+(0 \text{ or } 2^{++})$$

Mass  $m = 3918.4 \pm 1.9$  MeVFull width  $\Gamma = 20 \pm 5$  MeV (S = 1.1) **$X(3915)$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\omega J/\psi$	seen	222
$\pi^+ \pi^- \eta_c(1S)$	not seen	785
$\eta_c \eta$	not seen	665
$\eta_c \pi^0$	not seen	815
$K \bar{K}$	not seen	1896
$\gamma \gamma$	seen	1959

 **$\chi_{c2}(2P)$** 

$$J^G(J^{PC}) = 0^+(2^{++})$$

Mass  $m = 3927.2 \pm 2.6$  MeVFull width  $\Gamma = 24 \pm 6$  MeV

# Meson Summary Table

$\chi_{c2}(2P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\gamma\gamma$	seen	1964
$D\overline{D}$	seen	615
$D^+D^-$	seen	600
$D^0\overline{D}^0$	seen	615
$\pi^+\pi^-\eta_c(1S)$	not seen	792
$K\overline{K}$	not seen	1901

**X(4020)**

$I(J^P) = 1(?^?)$

Mass  $m = 4024.1 \pm 1.9$  MeV  
Full width  $\Gamma = 13 \pm 5$  MeV    ( $S = 1.7$ )

<b>X(4020) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$h_c(1P)\pi$	seen	450
$D^*\overline{D}^*$	seen	85
$D\overline{D}^* + c.c.$	not seen	542
$\eta_c\pi^+\pi^-$	not seen	872

**$\psi(4040)$**  [yyaa]

$I^G(J^{PC}) = 0^-(1^{--})$

Mass  $m = 4039 \pm 1$  MeV  
Full width  $\Gamma = 80 \pm 10$  MeV  
 $\Gamma_{ee} = 0.86 \pm 0.07$  keV  
 $\Gamma_{ee} < 2.9$  eV, CL = 90%  
 $\Gamma_{ee} < 4.6$  eV, CL = 90%

Due to the complexity of the  $c\overline{c}$  threshold region, in this listing, “seen” (“not seen”) means that a cross section for the mode in question has been measured at effective  $\sqrt{s}$  near this particle’s central mass value, more (less) than  $2\sigma$  above zero, without regard to any peaking behavior in  $\sqrt{s}$  or absence thereof. See mode listing(s) for details and references.

<b><math>\psi(4040)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$e^+e^-$	$(1.07 \pm 0.16) \times 10^{-5}$		2019
$D\overline{D}$	seen		775
$D^0\overline{D}^0$	seen		775
$D^+D^-$	seen		764
$D^*\overline{D} + c.c.$	seen		569
$D^*(2007)^0\overline{D}^0 + c.c.$	seen		575
$D^*(2010)^+D^- + c.c.$	seen		561
$D^*\overline{D}^*$	seen		193
$D^*(2007)^0\overline{D}^*(2007)^0$	seen		226
$D^*(2010)^+D^*(2010)^-$	seen		193
$D^0D^-\pi^+ + c.c.$ (excl. $D^*(2007)^0\overline{D}^0 + c.c.,$ $D^*(2010)^+D^- + c.c.)$	not seen		—
$D\overline{D}^*\pi$ (excl. $D^*\overline{D}^*$ )	not seen		—
$D^0\overline{D}^*\pi^- + c.c.$ (excl. $D^*(2010)^+D^*(2010)^-$ )	seen		—
$D_s^+D_s^-$	seen		452
$J/\psi\pi^+\pi^-$	$< 4$	$\times 10^{-3}$	90% 794
$J/\psi\pi^0\pi^0$	$< 2$	$\times 10^{-3}$	90% 797
$J/\psi\eta$	$(5.2 \pm 0.7)$	$\times 10^{-3}$	675
$J/\psi\pi^0$	$< 2.8$	$\times 10^{-4}$	90% 823
$J/\psi\pi^+\pi^-\pi^0$	$< 2$	$\times 10^{-3}$	90% 746
$\chi_{c1}\gamma$	$< 3.4$	$\times 10^{-3}$	90% 494
$\chi_{c2}\gamma$	$< 5$	$\times 10^{-3}$	90% 454
$\chi_{c1}\pi^+\pi^-\pi^0$	$< 1.1$	%	90% 306
$\chi_{c2}\pi^+\pi^-\pi^0$	$< 3.2$	%	90% 233
$h_c(1P)\pi^+\pi^-$	$< 3$	$\times 10^{-3}$	90% 403
$\phi\pi^+\pi^-$	$< 3$	$\times 10^{-3}$	90% 1880
$\Lambda\overline{\Lambda}\pi^+\pi^-$	$< 2.9$	$\times 10^{-4}$	90% 1578
$\Lambda\overline{\Lambda}\pi^0$	$< 9$	$\times 10^{-5}$	90% 1636
$\Lambda\overline{\Lambda}\eta$	$< 3.0$	$\times 10^{-4}$	90% 1452
$\Sigma^+\Sigma^-$	$< 1.3$	$\times 10^{-4}$	90% 1632
$\Sigma^0\overline{\Sigma}^0$	$< 7$	$\times 10^{-5}$	90% 1630
$\Xi^+\Xi^-$	$< 1.6$	$\times 10^{-4}$	90% 1527
$\Xi^0\overline{\Xi}^0$	$< 1.8$	$\times 10^{-4}$	90% 1533

**X(4140)**

$I^G(J^{PC}) = 0^+(?^?+)$

Mass  $m = 4146.9 \pm 3.1$  MeV    ( $S = 1.3$ )  
Full width  $\Gamma = 15^{+6}_{-5}$  MeV

<b>X(4140) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$J/\psi\phi$	seen	217
$\gamma\gamma$	not seen	2073

**$\psi(4160)$**  [yyaa]

$I^G(J^{PC}) = 0^-(1^{--})$

Mass  $m = 4191 \pm 5$  MeV  
Full width  $\Gamma = 70 \pm 10$  MeV  
 $\Gamma_{ee} = 0.48 \pm 0.22$  keV  
 $\Gamma_{ee} < 2.2$  eV, CL = 90%  
 $\Gamma_{ee}$

Due to the complexity of the  $c\overline{c}$  threshold region, in this listing, “seen” (“not seen”) means that a cross section for the mode in question has been measured at effective  $\sqrt{s}$  near this particle’s central mass value, more (less) than  $2\sigma$  above zero, without regard to any peaking behavior in  $\sqrt{s}$  or absence thereof. See mode listing(s) for details and references.

<b><math>\psi(4160)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$e^+e^-$	$(6.9 \pm 3.3) \times 10^{-6}$		2096
$\mu^+\mu^-$	seen		2093
$D\overline{D}$	seen		956
$D^0\overline{D}^0$	seen		956
$D^+D^-$	seen		947
$D^*\overline{D} + c.c.$	seen		798
$D^*(2007)^0\overline{D}^0 + c.c.$	seen		802
$D^*(2010)^+D^- + c.c.$	seen		792
$D^*\overline{D}^*$	seen		592
$D^*(2007)^0\overline{D}^*(2007)^0$	seen		604
$D^*(2010)^+D^*(2010)^-$	seen		592
$D^0D^-\pi^+ + c.c.$ (excl. $D^*(2007)^0\overline{D}^0 + c.c.,$ $D^*(2010)^+D^- + c.c.)$	not seen		—
$D\overline{D}^*\pi + c.c.$ (excl. $D^*\overline{D}^*$ )	seen		—
$D^0D^*\pi^- + c.c.$ (excl. $D^*(2010)^+D^*(2010)^-$ )	not seen		—
$D_s^+D_s^-$	not seen		720
$D_s^{*+}D_s^- + c.c.$	seen		385
$J/\psi\pi^+\pi^-$	$< 3$	$\times 10^{-3}$	90% 919
$J/\psi\pi^0\pi^0$	$< 3$	$\times 10^{-3}$	90% 922
$J/\psi K^+K^-$	$< 2$	$\times 10^{-3}$	90% 407
$J/\psi\eta$	$< 8$	$\times 10^{-3}$	90% 822
$J/\psi\pi^0$	$< 1$	$\times 10^{-3}$	90% 944
$J/\psi\eta'$	$< 5$	$\times 10^{-3}$	90% 457
$J/\psi\pi^+\pi^-\pi^0$	$< 1$	$\times 10^{-3}$	90% 879
$\psi(2S)\pi^+\pi^-$	$< 4$	$\times 10^{-3}$	90% 396
$\chi_{c1}\gamma$	$< 5$	$\times 10^{-3}$	90% 625
$\chi_{c2}\gamma$	$< 1.3$	%	90% 587
$\chi_{c1}\pi^+\pi^-\pi^0$	$< 2$	$\times 10^{-3}$	90% 496
$\chi_{c2}\pi^+\pi^-\pi^0$	$< 8$	$\times 10^{-3}$	90% 445
$h_c(1P)\pi^+\pi^-$	$< 5$	$\times 10^{-3}$	90% 556
$h_c(1P)\pi^0\pi^0$	$< 2$	$\times 10^{-3}$	90% 560
$h_c(1P)\eta$	$< 2$	$\times 10^{-3}$	90% 348
$h_c(1P)\pi^0$	$< 4$	$\times 10^{-4}$	90% 600
$\phi\pi^+\pi^-$	$< 2$	$\times 10^{-3}$	90% 1961
$\gamma X(3872) \rightarrow \gamma J/\psi\pi^+\pi^-$	$< 6.8$	$\times 10^{-5}$	90% —
$\gamma X(3915) \rightarrow \gamma J/\psi\pi^+\pi^-$	$< 1.36$	$\times 10^{-4}$	90% —
$\gamma X(3930) \rightarrow \gamma J/\psi\pi^+\pi^-$	$< 1.18$	$\times 10^{-4}$	90% —
$\gamma X(3940) \rightarrow \gamma J/\psi\pi^+\pi^-$	$< 1.47$	$\times 10^{-4}$	90% —
$\gamma X(3872) \rightarrow \gamma\gamma J/\psi$	$< 1.05$	$\times 10^{-4}$	90% —
$\gamma X(3915) \rightarrow \gamma\gamma J/\psi$	$< 1.26$	$\times 10^{-4}$	90% —
$\gamma X(3930) \rightarrow \gamma\gamma J/\psi$	$< 8.8$	$\times 10^{-5}$	90% —
$\gamma X(3940) \rightarrow \gamma\gamma J/\psi$	$< 1.79$	$\times 10^{-4}$	90% —

**X(4260)**

$I^G(J^{PC}) = ?^?(1^{--})$

Mass  $m = 4251 \pm 9$  MeV    ( $S = 1.6$ )  
Full width  $\Gamma = 120 \pm 12$  MeV    ( $S = 1.1$ )

Meson Summary Table

X(4260) DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$J/\psi \pi^+ \pi^-$	seen	967
$J/\psi f_0(980), f_0(980) \rightarrow \pi^+ \pi^-$	seen	—
$X(3900)^\pm \pi^\mp, X^\pm \rightarrow J/\psi \pi^\pm$	seen	—
$J/\psi \pi^0 \pi^0$	seen	969
$J/\psi K^+ K^-$	seen	512
$J/\psi K_S^0 K_S^0$	not seen	501
$X(3872) \gamma$	seen	363
$J/\psi \eta$	not seen	876
$J/\psi \pi^0$	not seen	991
$J/\psi \eta'$	not seen	552
$J/\psi \pi^+ \pi^- \pi^0$	not seen	930
$J/\psi \eta \pi^0$	not seen	801
$J/\psi \eta \eta$	not seen	311
$\psi(2S) \pi^+ \pi^-$	not seen	459
$\psi(2S) \eta$	not seen	129
$\chi_{c0} \omega$	not seen	265
$\chi_{c1} \gamma$	not seen	676
$\chi_{c2} \gamma$	not seen	638
$\chi_{c1} \pi^+ \pi^- \pi^0$	not seen	560
$\chi_{c2} \pi^+ \pi^- \pi^0$	not seen	512
$h_c(1P) \pi^+ \pi^-$	not seen	613
$\phi \pi^+ \pi^-$	not seen	1993
$\phi f_0(980) \rightarrow \phi \pi^+ \pi^-$	not seen	—
$D \bar{D}$	not seen	1020
$D^0 \bar{D}^0$	not seen	1020
$D^+ D^-$	not seen	1011
$D^* \bar{D} + c.c.$	not seen	887
$D^*(2007)^0 \bar{D}^0 + c.c.$	not seen	—
$D^*(2010)^+ D^- + c.c.$	not seen	—
$D^* \bar{D}^*$	not seen	691
$D^*(2007)^0 \bar{D}^*(2007)^0$	not seen	701
$D^*(2010)^+ D^*(2010)^-$	not seen	691
$D^0 D^- \pi^+ + c.c. (excl.$	not seen	—
$D^*(2007)^0 \bar{D}^{*0} + c.c.,$		
$D^*(2010)^+ D^- + c.c.)$		
$D \bar{D}^* \pi + c.c. (excl. D^* \bar{D}^*)$	not seen	723
$D^0 D^{*-} \pi^+ + c.c. (excl.$	not seen	—
$D^*(2010)^+ D^*(2010)^-$		
$D^0 D^*(2010)^- \pi^+ + c.c.$	not seen	716
$D^* \bar{D}^* \pi$	not seen	448
$D_s^+ D_s^-$	not seen	803
$D_s^{*+} D_s^- + c.c.$	not seen	615
$D_s^{*+} D_s^{*-}$	not seen	239
$\rho \bar{\rho}$	not seen	1907
$K_S^0 K^\pm \pi^\mp$	not seen	2048
$K^+ K^- \pi^0$	not seen	2049

X(4360)

$I^G(J^{PC}) = ??(1^{--})$

Mass  $m = 4346 \pm 6$  MeV  
X(4360) WIDTH =  $102 \pm 10$  MeV  
 $\Gamma_{ee}$   
 $\Gamma_{ee} < 0.57$  eV, CL = 90%  
 $\Gamma_{ee} < 1.9$  eV, CL = 90%

X(4360) DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\psi(2S) \pi^+ \pi^-$	seen	552
$\psi(3823) \pi^+ \pi^-$	possibly seen	416

$\psi(4415)$  [yyaa]

$I^G(J^{PC}) = 0^-(1^{--})$

Mass  $m = 4421 \pm 4$  MeV  
Full width  $\Gamma = 62 \pm 20$  MeV  
 $\Gamma_{ee} = 0.58 \pm 0.07$  keV  
 $\Gamma_{ee} < 3.6$  eV, CL = 90%  
 $\Gamma_{ee} < 0.47$  eV, CL = 90%  
 $\Gamma_{ee} < 2.3$  eV, CL = 90%

Due to the complexity of the  $c\bar{c}$  threshold region, in this listing, “seen” (“not seen”) means that a cross section for the mode in question has been measured at effective  $\sqrt{s}$  near this particle’s central mass value, more (less) than  $2\sigma$  above zero, without regard to any peaking behavior in  $\sqrt{s}$  or absence thereof. See mode listing(s) for details and references.

$\psi(4415)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$D \bar{D}$	seen		1187
$D^0 \bar{D}^0$	seen		1187
$D^+ D^-$	seen		1179
$D^* \bar{D} + c.c.$	seen		1063
$D^*(2007)^0 \bar{D}^0 + c.c.$	seen		1067
$D^*(2010)^+ D^- + c.c.$	seen		1059
$D^* \bar{D}^*$	seen		919
$D^*(2007)^0 \bar{D}^*(2007)^0 + c.c.$	seen		927
$D^*(2010)^+ D^*(2010)^- + c.c.$	seen		919
$D^0 D^- \pi^+ (excl. D^*(2007)^0 \bar{D}^0$	< 2.3 %	90%	—
$+ c.c., D^*(2010)^+ D^- + c.c.)$			
$D \bar{D}_2^*(2460) \rightarrow D^0 D^- \pi^+ + c.c.$	(10 $\pm$ 4 ) %		—
$D^0 D^{*-} \pi^+ + c.c.$	< 11 %	90%	926
$D_s^+ D_s^-$	not seen		1006
$\omega \chi_{c2}$	possibly seen		330
$D_s^{*+} D_s^- + c.c.$	seen		—
$D_s^{*+} D_s^{*-}$	not seen		652
$\psi(3823) \pi^+ \pi^-$	possibly seen		494
$J/\psi \eta$	< 6 $\times 10^{-3}$	90%	1022
$\chi_{c1} \gamma$	< 8 $\times 10^{-4}$	90%	817
$\chi_{c2} \gamma$	< 4 $\times 10^{-3}$	90%	780
$e^+ e^-$	( 9.4 $\pm$ 3.2 ) $\times 10^{-6}$		2210

X(4430)<sup>±</sup>

$I(J^P) = ?(1^+)$

Quantum numbers not established.

Mass  $m = 4478^{+15}_{-18}$  MeV  
Full width  $\Gamma = 181 \pm 31$  MeV

X(4430) <sup>±</sup> DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\pi^+ \psi(2S)$	seen	711
$\pi^+ J/\psi$	seen	1162

X(4660)

$I^G(J^{PC}) = ??(1^{--})$

X(4660) MASS =  $4643 \pm 9$  MeV (S = 1.2)  
X(4660) WIDTH =  $72 \pm 11$  MeV  
 $\Gamma_{ee}$   
 $\Gamma_{ee} < 0.45$  eV, CL = 90%  
 $\Gamma_{ee} < 2.1$  eV, CL = 90%

X(4660) DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\psi(2S) \pi^+ \pi^-$	seen	820

$b\bar{b}$  MESONS

$\eta_b(1S)$

$I^G(J^{PC}) = 0^+(0^{--})$

Mass  $m = 9399.0 \pm 2.3$  MeV (S = 1.6)  
Full width  $\Gamma = 10^{+5}_{-4}$  MeV

$\eta_b(1S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
hadrons	seen		—
$3h^+ 3h^-$	not seen		4673
$2h^+ 2h^-$	not seen		4689
$\gamma \gamma$	not seen		4700
$\mu^+ \mu^-$	< 9 $\times 10^{-3}$	90%	4698
$\tau^+ \tau^-$	< 8 %	90%	4351

# Meson Summary Table

<div><div><math>\tau(1S)</math></div><div>Mass <math>m = 9460.30 \pm 0.26</math> MeV (S = 3.3) Full width <math>\Gamma = 54.02 \pm 1.25</math> keV <math>\Gamma_{ee} = 1.340 \pm 0.018</math> keV</div></div>	$I^G(J^{PC}) = 0^-(1^--)$			
	<b><math>\tau(1S)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\frac{p}{(\text{MeV}/c)}$
	$\tau^+ \tau^-$	( 2.60 $\pm$ 0.10 ) %		4384
	$e^+ e^-$	( 2.38 $\pm$ 0.11 ) %		4730
	$\mu^+ \mu^-$	( 2.48 $\pm$ 0.05 ) %		4729
Hadronic decays				
$g g g$	(81.7 $\pm$ 0.7 ) %		—	
$\gamma g g$	( 2.2 $\pm$ 0.6 ) %		—	
$\eta'(958)$ anything	( 2.94 $\pm$ 0.24 ) %		—	
$J/\psi(1S)$ anything	( 6.5 $\pm$ 0.7 ) $\times 10^{-4}$		4223	
$J/\psi(1S) \eta_c$	< 2.2 $\times 10^{-6}$	90%	3623	
$J/\psi(1S) \chi_{c0}$	< 3.4 $\times 10^{-6}$	90%	3429	
$J/\psi(1S) \chi_{c1}$	( 3.9 $\pm$ 1.2 ) $\times 10^{-6}$		3382	
$J/\psi(1S) \chi_{c2}$	< 1.4 $\times 10^{-6}$	90%	3359	
$J/\psi(1S) \eta_c(2S)$	< 2.2 $\times 10^{-6}$	90%	3316	
$J/\psi(1S) X(3940)$	< 5.4 $\times 10^{-6}$	90%	3148	
$J/\psi(1S) X(4160)$	< 5.4 $\times 10^{-6}$	90%	3018	
$\chi_{c0}$ anything	< 5 $\times 10^{-3}$	90%	—	
$\chi_{c1}$ anything	( 2.3 $\pm$ 0.7 ) $\times 10^{-4}$		—	
$\chi_{c2}$ anything	( 3.4 $\pm$ 1.0 ) $\times 10^{-4}$		—	
$\psi(2S)$ anything	( 2.7 $\pm$ 0.9 ) $\times 10^{-4}$		—	
$\psi(2S) \eta_c$	< 3.6 $\times 10^{-6}$	90%	3345	
$\psi(2S) \chi_{c0}$	< 6.5 $\times 10^{-6}$	90%	3124	
$\psi(2S) \chi_{c1}$	< 4.5 $\times 10^{-6}$	90%	3070	
$\psi(2S) \chi_{c2}$	< 2.1 $\times 10^{-6}$	90%	3043	
$\psi(2S) \eta_c(2S)$	< 3.2 $\times 10^{-6}$	90%	2993	
$\psi(2S) X(3940)$	< 2.9 $\times 10^{-6}$	90%	2797	
$\psi(2S) X(4160)$	< 2.9 $\times 10^{-6}$	90%	2642	
$\rho \pi$	< 3.68 $\times 10^{-6}$	90%	4697	
$\omega \pi^0$	< 3.90 $\times 10^{-6}$	90%	4697	
$\pi^+ \pi^-$	< 5 $\times 10^{-4}$	90%	4728	
$K^+ K^-$	< 5 $\times 10^{-4}$	90%	4704	
$\rho \overline{\rho}$	< 5 $\times 10^{-4}$	90%	4636	
$\pi^+ \pi^- \pi^0$	( 2.1 $\pm$ 0.8 ) $\times 10^{-6}$		4725	
$\phi K^+ K^-$	( 2.4 $\pm$ 0.5 ) $\times 10^{-6}$		4622	
$\omega \pi^+ \pi^-$	( 4.5 $\pm$ 1.0 ) $\times 10^{-6}$		4694	
$K^*(892)^0 K^- \pi^+ + \text{c.c.}$	( 4.4 $\pm$ 0.8 ) $\times 10^{-6}$		4667	
$\phi f_2'(1525)$	< 1.63 $\times 10^{-6}$	90%	4549	
$\omega f_2'(1270)$	< 1.79 $\times 10^{-6}$	90%	4611	
$\rho(770) a_2(1320)$	< 2.24 $\times 10^{-6}$	90%	4605	
$K^*(892)^0 \overline{K}_2^0(1430)^0 + \text{c.c.}$	( 3.0 $\pm$ 0.8 ) $\times 10^{-6}$		4579	
$K_1(1270)^\pm K^\mp$	< 2.41 $\times 10^{-6}$	90%	4631	
$K_1(1400)^\pm K^\mp$	( 1.0 $\pm$ 0.4 ) $\times 10^{-6}$		4613	
$b_1(1235)^\pm \pi^\mp$	< 1.25 $\times 10^{-6}$	90%	4649	
$\pi^+ \pi^- \pi^0 \pi^0$	( 1.28 $\pm$ 0.30 ) $\times 10^{-5}$		4720	
$K_S^0 K^+ \pi^- + \text{c.c.}$	( 1.6 $\pm$ 0.4 ) $\times 10^{-6}$		4696	
$K^*(892)^0 \overline{K}^0 + \text{c.c.}$	( 2.9 $\pm$ 0.9 ) $\times 10^{-6}$		4675	
$K^*(892)^- K^+ + \text{c.c.}$	< 1.11 $\times 10^{-6}$	90%	4675	
$D^*(2010)^\pm$ anything	( 2.52 $\pm$ 0.20 ) %		—	
$^2H$ anything	( 2.85 $\pm$ 0.25 ) $\times 10^{-5}$		—	
Sum of 100 exclusive modes	( 1.200 $\pm$ 0.017 ) %		—	
Radiative decays				
$\gamma \pi^+ \pi^-$	( 6.3 $\pm$ 1.8 ) $\times 10^{-5}$		4728	
$\gamma \pi^0 \pi^0$	( 1.7 $\pm$ 0.7 ) $\times 10^{-5}$		4728	
$\gamma \pi^0 \eta$	< 2.4 $\times 10^{-6}$	90%	4713	
$\gamma K^+ K^-$	[zzaa] ( 1.14 $\pm$ 0.13 ) $\times 10^{-5}$		4704	
$\gamma \rho \overline{\rho}$	[aabb] < 6 $\times 10^{-6}$	90%	4636	
$\gamma 2h^+ 2h^-$	( 7.0 $\pm$ 1.5 ) $\times 10^{-4}$		4720	
$\gamma 3h^+ 3h^-$	( 5.4 $\pm$ 2.0 ) $\times 10^{-4}$		4703	
$\gamma 4h^+ 4h^-$	( 7.4 $\pm$ 3.5 ) $\times 10^{-4}$		4679	
$\gamma \pi^+ \pi^- K^+ K^-$	( 2.9 $\pm$ 0.9 ) $\times 10^{-4}$		4686	
$\gamma 2\pi^+ 2\pi^-$	( 2.5 $\pm$ 0.9 ) $\times 10^{-4}$		4720	
$\gamma 3\pi^+ 3\pi^-$	( 2.5 $\pm$ 1.2 ) $\times 10^{-4}$		4703	
$\gamma 2\pi^+ 2\pi^- K^+ K^-$	( 2.4 $\pm$ 1.2 ) $\times 10^{-4}$		4658	
$\gamma \pi^+ \pi^- \rho \overline{\rho}$	( 1.5 $\pm$ 0.6 ) $\times 10^{-4}$		4604	
$\gamma 2\pi^+ 2\pi^- \rho \overline{\rho}$	( 4 $\pm$ 6 ) $\times 10^{-5}$		4563	
$\gamma 2K^+ 2K^-$	( 2.0 $\pm$ 2.0 ) $\times 10^{-5}$		4601	
$\gamma \eta'(958)$	< 1.9 $\times 10^{-6}$	90%	4682	
$\gamma \eta$	< 1.0 $\times 10^{-6}$	90%	4714	
$\gamma f_0(980)$	< 3 $\times 10^{-5}$	90%	4678	
$\gamma f_2'(1525)$	( 3.8 $\pm$ 0.9 ) $\times 10^{-5}$		4607	
$\gamma f_2'(1270)$	( 1.01 $\pm$ 0.09 ) $\times 10^{-4}$		4644	
$\gamma \eta(1405)$	< 8.2 $\times 10^{-5}$	90%	4625	
$\gamma f_0(1500)$	< 1.5 $\times 10^{-5}$	90%	4611	
$\gamma f_0(1710)$	< 2.6 $\times 10^{-4}$	90%	4573	
$\gamma f_0(1710) \rightarrow \gamma K^+ K^-$	< 7 $\times 10^{-6}$	90%	—	
$\gamma f_0(1710) \rightarrow \gamma \pi^0 \pi^0$	< 1.4 $\times 10^{-6}$	90%	—	
$\gamma f_0(1710) \rightarrow \gamma \eta \eta$	< 1.8 $\times 10^{-6}$	90%	—	
$\gamma f_4(2050)$	< 5.3 $\times 10^{-5}$	90%	4515	
$\gamma f_0(2200) \rightarrow \gamma K^+ K^-$	< 2 $\times 10^{-4}$	90%	4475	
$\gamma f_J(2220) \rightarrow \gamma K^+ K^-$	< 8 $\times 10^{-7}$	90%	4469	
$\gamma f_J(2220) \rightarrow \gamma \pi^+ \pi^-$	< 6 $\times 10^{-7}$	90%	—	
$\gamma f_J(2220) \rightarrow \gamma \rho \overline{\rho}$	< 1.1 $\times 10^{-6}$	90%	—	
$\gamma \eta(2225) \rightarrow \gamma \phi \phi$	< 3 $\times 10^{-3}$	90%	4469	
$\gamma \eta_c(1S)$	< 5.7 $\times 10^{-5}$	90%	4260	
$\gamma \chi_{c0}$	< 6.5 $\times 10^{-4}$	90%	4114	
$\gamma \chi_{c1}$	< 2.3 $\times 10^{-5}$	90%	4079	
$\gamma \chi_{c2}$	< 7.6 $\times 10^{-6}$	90%	4062	
$\gamma X(3872) \rightarrow \pi^+ \pi^- J/\psi$	< 1.6 $\times 10^{-6}$	90%	—	
$\gamma X(3872) \rightarrow \pi^+ \pi^- \pi^0 J/\psi$	< 2.8 $\times 10^{-6}$	90%	—	
$\gamma X(3915) \rightarrow \omega J/\psi$	< 3.0 $\times 10^{-6}$	90%	—	
$\gamma X(4140) \rightarrow \phi J/\psi$	< 2.2 $\times 10^{-6}$	90%	—	
$\gamma X$	[bbbb] < 4.5 $\times 10^{-6}$	90%	—	
$\gamma X \overline{X} (m_X < 3.1 \text{ GeV})$	[ccbb] < 1 $\times 10^{-3}$	90%	—	
$\gamma X \overline{X} (m_X < 4.5 \text{ GeV})$	[ddbb] < 2.4 $\times 10^{-4}$	90%	—	
$\gamma X \rightarrow \gamma + \geq 4 \text{ prongs}$	[eebb] < 1.78 $\times 10^{-4}$	95%	—	
$\gamma a_1^0 \rightarrow \gamma \mu^+ \mu^-$	[ffbb] < 9 $\times 10^{-6}$	90%	—	
$\gamma a_1^0 \rightarrow \gamma \tau^+ \tau^-$	[zzaa] < 1.30 $\times 10^{-4}$	90%	—	
$\gamma a_1^0 \rightarrow \gamma g g$	[ggbb] < 1 %	90%	—	
$\gamma a_1^0 \rightarrow \gamma s \overline{s}$	[ggbb] < 1 $\times 10^{-3}$	90%	—	
Lepton Family number (LF) violating modes				
$\mu^\pm \tau^\mp$	LF < 6.0 $\times 10^{-6}$	95%	4563	
Other decays				
invisible	< 3.0 $\times 10^{-4}$	90%	—	
<div><div><math>\chi_{b0}(1P)</math> [hhbb]</div><div><math>I^G(J^{PC}) = 0^+(0^+ +)</math> <math>J</math> needs confirmation. Mass <math>m = 9859.44 \pm 0.42 \pm 0.31</math> MeV</div></div>				
<b><math>\chi_{b0}(1P)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\frac{p}{(\text{MeV}/c)}$	
$\gamma T(1S)$	( 1.76 $\pm$ 0.35 ) %		391	
$D^0 X$	< 10.4 %	90%	—	
$\pi^+ \pi^- K^+ K^- \pi^0$	< 1.6 $\times 10^{-4}$	90%	4875	
$2\pi^+ \pi^- K^- K_S^0$	< 5 $\times 10^{-5}$	90%	4875	
$2\pi^+ \pi^- K^- K_S^0 2\pi^0$	< 5 $\times 10^{-4}$	90%	4846	
$2\pi^+ 2\pi^- 2\pi^0$	< 2.1 $\times 10^{-4}$	90%	4905	
$2\pi^+ 2\pi^- K^+ K^-$	( 1.1 $\pm$ 0.6 ) $\times 10^{-4}$		4861	
$2\pi^+ 2\pi^- K^+ K^- \pi^0$	< 2.7 $\times 10^{-4}$	90%	4846	
$2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	< 5 $\times 10^{-4}$	90%	4828	
$3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	< 1.6 $\times 10^{-4}$	90%	4827	
$3\pi^+ 3\pi^-$	< 8 $\times 10^{-5}$	90%	4904	
$3\pi^+ 3\pi^- 2\pi^0$	< 6 $\times 10^{-4}$	90%	4881	
$3\pi^+ 3\pi^- K^+ K^-$	( 2.4 $\pm$ 1.2 ) $\times 10^{-4}$		4827	
$3\pi^+ 3\pi^- K^+ K^- \pi^0$	< 1.0 $\times 10^{-3}$	90%	4808	
$4\pi^+ 4\pi^-$	< 8 $\times 10^{-5}$	90%	4880	
$4\pi^+ 4\pi^- 2\pi^0$	< 2.1 $\times 10^{-3}$	90%	4850	
$J/\psi J/\psi$	< 7 $\times 10^{-5}$	90%	3836	
$J/\psi \psi(2S)$	< 1.2 $\times 10^{-4}$	90%	3571	
$\psi(2S) \psi(2S)$	< 3.1 $\times 10^{-5}$	90%	3273	
<div><div><math>\chi_{b1}(1P)</math> [hhbb]</div><div><math>I^G(J^{PC}) = 0^+(1^+ +)</math> <math>J</math> needs confirmation. Mass <math>m = 9892.78 \pm 0.26 \pm 0.31</math> MeV</div></div>				



## Meson Summary Table

$\chi_{b1}(1P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\gamma T(1S)$	(33.9 $\pm$ 2.2) %		423
$D^0 X$	(12.6 $\pm$ 2.2) %		—
$\pi^+ \pi^- K^+ K^- \pi^0$	( 2.0 $\pm$ 0.6) $\times 10^{-4}$		4892
$2\pi^+ \pi^- K^- K_S^0$	( 1.3 $\pm$ 0.5) $\times 10^{-4}$		4892
$2\pi^+ \pi^- K^- K_S^0 2\pi^0$	< 6 $\times 10^{-4}$	90%	4863
$2\pi^+ 2\pi^- 2\pi^0$	( 8.0 $\pm$ 2.5) $\times 10^{-4}$		4921
$2\pi^+ 2\pi^- K^+ K^-$	( 1.5 $\pm$ 0.5) $\times 10^{-4}$		4878
$2\pi^+ 2\pi^- K^+ K^- \pi^0$	( 3.5 $\pm$ 1.2) $\times 10^{-4}$		4863
$2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	( 8.6 $\pm$ 3.2) $\times 10^{-4}$		4845
$3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	( 9.3 $\pm$ 3.3) $\times 10^{-4}$		4844
$3\pi^+ 3\pi^-$	( 1.9 $\pm$ 0.6) $\times 10^{-4}$		4921
$3\pi^+ 3\pi^- 2\pi^0$	( 1.7 $\pm$ 0.5) $\times 10^{-3}$		4898
$3\pi^+ 3\pi^- K^+ K^-$	( 2.6 $\pm$ 0.8) $\times 10^{-4}$		4844
$3\pi^+ 3\pi^- K^+ K^- \pi^0$	( 7.5 $\pm$ 2.6) $\times 10^{-4}$		4825
$4\pi^+ 4\pi^-$	( 2.6 $\pm$ 0.9) $\times 10^{-4}$		4897
$4\pi^+ 4\pi^- 2\pi^0$	( 1.4 $\pm$ 0.6) $\times 10^{-3}$		4867
$J/\psi J/\psi$	< 2.7 $\times 10^{-5}$	90%	3857
$J/\psi \psi(2S)$	< 1.7 $\times 10^{-5}$	90%	3594
$\psi(2S) \psi(2S)$	< 6 $\times 10^{-5}$	90%	3298

<b><math>h_b(1P)</math></b>	$J^G(J^{PC}) = ?^?(1^+ -)$
Mass $m = 9899.3 \pm 0.8$ MeV	
$h_b(1P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ ) $p$ (MeV/c)
$\eta_b(1S) \gamma$	(52 $^{+6}_{-5}$ ) % 488

<b><math>\chi_{b2}(1P)</math> <math>[hhbb]</math></b>	$J^G(J^{PC}) = 0^+(2^+ +)$ $J$ needs confirmation.
Mass $m = 9912.21 \pm 0.26 \pm 0.31$ MeV	

$\chi_{b2}(1P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\gamma T(1S)$	(19.1 $\pm$ 1.2) %		442
$D^0 X$	< 7.9 %	90%	—
$\pi^+ \pi^- K^+ K^- \pi^0$	( 8 $\pm$ 5 ) $\times 10^{-5}$		4902
$2\pi^+ \pi^- K^- K_S^0$	< 1.0 $\times 10^{-4}$	90%	4901
$2\pi^+ \pi^- K^- K_S^0 2\pi^0$	( 5.3 $\pm$ 2.4) $\times 10^{-4}$		4873
$2\pi^+ 2\pi^- 2\pi^0$	( 3.5 $\pm$ 1.4) $\times 10^{-4}$		4931
$2\pi^+ 2\pi^- K^+ K^-$	( 1.1 $\pm$ 0.4) $\times 10^{-4}$		4888
$2\pi^+ 2\pi^- K^+ K^- \pi^0$	( 2.1 $\pm$ 0.9) $\times 10^{-4}$		4872
$2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	( 3.9 $\pm$ 1.8) $\times 10^{-4}$		4855
$3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	< 5 $\times 10^{-4}$	90%	4854
$3\pi^+ 3\pi^-$	( 7.0 $\pm$ 3.1) $\times 10^{-5}$		4931
$3\pi^+ 3\pi^- 2\pi^0$	( 1.0 $\pm$ 0.4) $\times 10^{-3}$		4908
$3\pi^+ 3\pi^- K^+ K^-$	< 8 $\times 10^{-5}$	90%	4854
$3\pi^+ 3\pi^- K^+ K^- \pi^0$	( 3.6 $\pm$ 1.5) $\times 10^{-4}$		4835
$4\pi^+ 4\pi^-$	( 8 $\pm$ 4 ) $\times 10^{-5}$		4907
$4\pi^+ 4\pi^- 2\pi^0$	( 1.8 $\pm$ 0.7) $\times 10^{-3}$		4877
$J/\psi J/\psi$	< 4 $\times 10^{-5}$	90%	3869
$J/\psi \psi(2S)$	< 5 $\times 10^{-5}$	90%	3608
$\psi(2S) \psi(2S)$	< 1.6 $\times 10^{-5}$	90%	3313

<b><math>T(2S)</math></b>	$J^G(J^{PC}) = 0^-(1^- -)$
Mass $m = 10023.26 \pm 0.31$ MeV	
$m_{T(3S)} - m_{T(2S)} = 331.50 \pm 0.13$ MeV	
Full width $\Gamma = 31.98 \pm 2.63$ keV	
$\Gamma_{ee} = 0.612 \pm 0.011$ keV	

$T(2S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$T(1S) \pi^+ \pi^-$	(17.85 $\pm$ 0.26) %		475
$T(1S) \pi^0 \pi^0$	( 8.6 $\pm$ 0.4 ) %		480
$\tau^+ \tau^-$	( 2.00 $\pm$ 0.21) %		4686
$\mu^+ \mu^-$	( 1.93 $\pm$ 0.17) %	S=2.2	5011
$e^+ e^-$	( 1.91 $\pm$ 0.16) %		5012
$T(1S) \pi^0$	< 4 $\times 10^{-5}$	CL=90%	531
$T(1S) \eta$	( 2.9 $\pm$ 0.4 ) $\times 10^{-4}$	S=2.0	126
$J/\psi(1S)$ anything	< 6 $\times 10^{-3}$	CL=90%	4533
$J/\psi(1S) \eta_c$	< 5.4 $\times 10^{-6}$	CL=90%	3984
$J/\psi(1S) \chi_{c0}$	< 3.4 $\times 10^{-6}$	CL=90%	3808
$J/\psi(1S) \chi_{c1}$	< 1.2 $\times 10^{-6}$	CL=90%	3765
$J/\psi(1S) \chi_{c2}$	< 2.0 $\times 10^{-6}$	CL=90%	3744
$J/\psi(1S) \eta_c(2S)$	< 2.5 $\times 10^{-6}$	CL=90%	3706
$J/\psi(1S) X(3940)$	< 2.0 $\times 10^{-6}$	CL=90%	3555
$J/\psi(1S) X(4160)$	< 2.0 $\times 10^{-6}$	CL=90%	3440
$\psi(2S) \eta_c$	< 5.1 $\times 10^{-6}$	CL=90%	3732
$\psi(2S) \chi_{c0}$	< 4.7 $\times 10^{-6}$	CL=90%	3536
$\psi(2S) \chi_{c1}$	< 2.5 $\times 10^{-6}$	CL=90%	3488
$\psi(2S) \chi_{c2}$	< 1.9 $\times 10^{-6}$	CL=90%	3464
$\psi(2S) \eta_c(2S)$	< 3.3 $\times 10^{-6}$	CL=90%	3421
$\psi(2S) X(3940)$	< 3.9 $\times 10^{-6}$	CL=90%	3250
$\psi(2S) X(4160)$	< 3.9 $\times 10^{-6}$	CL=90%	3118
$^2\overline{H}$ anything	( 2.78 $^{+0.30}_{-0.26}$ ) $\times 10^{-5}$	S=1.2	—
hadrons	(94 $\pm$ 11 ) %		—
$ggg$	(58.8 $\pm$ 1.2 ) %		—
$\gamma g g$	( 1.87 $\pm$ 0.28) %		—
$\phi K^+ K^-$	( 1.6 $\pm$ 0.4 ) $\times 10^{-6}$		4910
$\omega \pi^+ \pi^-$	< 2.58 $\times 10^{-6}$	CL=90%	4977
$K^*(892)^0 K^- \pi^+ +$ c.c.	( 2.3 $\pm$ 0.7 ) $\times 10^{-6}$		4952
$\phi f_2'(1525)$	< 1.33 $\times 10^{-6}$	CL=90%	4841
$\omega f_2(1270)$	< 5.7 $\times 10^{-7}$	CL=90%	4899
$\rho(770) a_2(1320)$	< 8.8 $\times 10^{-7}$	CL=90%	4894
$K^*(892)^0 \overline{K}_2^*(1430)^0 +$ c.c.	( 1.5 $\pm$ 0.6 ) $\times 10^{-6}$		4869
$K_1(1270)^\pm K^\mp$	< 3.22 $\times 10^{-6}$	CL=90%	4918
$K_1^*(1400)^\pm K^\mp$	< 8.3 $\times 10^{-7}$	CL=90%	4901
$b_1(1235)^\pm \pi^\mp$	< 4.0 $\times 10^{-7}$	CL=90%	4935
$\rho \pi$	< 1.16 $\times 10^{-6}$	CL=90%	4981
$\pi^+ \pi^- \pi^0$	< 8.0 $\times 10^{-7}$	CL=90%	5007
$\omega \pi^0$	< 1.63 $\times 10^{-6}$	CL=90%	4980
$\pi^+ \pi^- \pi^0 \pi^0$	( 1.30 $\pm$ 0.28) $\times 10^{-5}$		5002
$K_S^0 K^+ \pi^- +$ c.c.	( 1.14 $\pm$ 0.33) $\times 10^{-6}$		4979
$K^*(892)^0 \overline{K}^0 +$ c.c.	< 4.22 $\times 10^{-6}$	CL=90%	4959
$K^*(892)^- K^+ +$ c.c.	< 1.45 $\times 10^{-6}$	CL=90%	4960
Sum of 100 exclusive modes	( 2.90 $\pm$ 0.30) $\times 10^{-3}$		—

## Radiative decays

$\gamma \chi_{b1}(1P)$	( 6.9 $\pm$ 0.4 ) %		130
$\gamma \chi_{b2}(1P)$	( 7.15 $\pm$ 0.35) %		110
$\gamma \chi_{b0}(1P)$	( 3.8 $\pm$ 0.4 ) %		162
$\gamma f_0(1710)$	< 5.9 $\times 10^{-4}$	CL=90%	4864
$\gamma f_2'(1525)$	< 5.3 $\times 10^{-4}$	CL=90%	4896
$\gamma f_2(1270)$	< 2.41 $\times 10^{-4}$	CL=90%	4930
$\gamma \eta_c(1S)$	< 2.7 $\times 10^{-5}$	CL=90%	4568
$\gamma \chi_{c0}$	< 1.0 $\times 10^{-4}$	CL=90%	4430
$\gamma \chi_{c1}$	< 3.6 $\times 10^{-6}$	CL=90%	4397
$\gamma \chi_{c2}$	< 1.5 $\times 10^{-5}$	CL=90%	4381
$\gamma X(3872) \rightarrow \pi^+ \pi^- J/\psi$	< 8 $\times 10^{-7}$	CL=90%	—
$\gamma X(3872) \rightarrow \pi^+ \pi^- \pi^0 J/\psi$	< 2.4 $\times 10^{-6}$	CL=90%	—
$\gamma X(3915) \rightarrow \omega J/\psi$	< 2.8 $\times 10^{-6}$	CL=90%	—
$\gamma X(4140) \rightarrow \phi J/\psi$	< 1.2 $\times 10^{-6}$	CL=90%	—
$\gamma X(4350) \rightarrow \phi J/\psi$	< 1.3 $\times 10^{-6}$	CL=90%	—
$\gamma \eta_b(1S)$	( 3.9 $\pm$ 1.5 ) $\times 10^{-4}$		605
$\gamma \eta_b(1S) \rightarrow \gamma$ Sum of 26 exclusive modes	< 3.7 $\times 10^{-6}$	CL=90%	—
$\gamma X_{b\overline{b}} \rightarrow \gamma$ Sum of 26 exclusive modes	< 4.9 $\times 10^{-6}$	CL=90%	—
$\gamma X \rightarrow \gamma + \geq 4$ prongs $[iibb]$	< 1.95 $\times 10^{-4}$	CL=95%	—
$\gamma A^0 \rightarrow \gamma$ hadrons	< 8 $\times 10^{-5}$	CL=90%	—
$\gamma a_1^0 \rightarrow \gamma \mu^+ \mu^-$	< 8.3 $\times 10^{-6}$	CL=90%	—

## Lepton Family number (LF) violating modes

$e^\pm \tau^\mp$	LF	< 3.2 $\times 10^{-6}$	CL=90%	4854
$\mu^\pm \tau^\mp$	LF	< 3.3 $\times 10^{-6}$	CL=90%	4854

# Meson Summary Table

<b><math>T(1D)</math></b>	$J^G(J^{PC}) = 0^-(2^{--})$	
Mass $m = 10163.7 \pm 1.4$ MeV    (S = 1.7)		
<b><math>T(1D)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\gamma\gamma\ T(1S)$	seen	679
$\gamma\chi_{bJ}(1P)$	seen	300
$\eta\ T(1S)$	not seen	426
$\pi^+\pi^-\ T(1S)$	$(6.6\pm1.6)\times10^{-3}$	623

$\chi_{b0}(2P)$  <sup>[hhbb]</sup>

$J^G(J^{PC}) = 0^+(0^{++})$   
 $J$  needs confirmation.

Mass  $m = 10232.5 \pm 0.4 \pm 0.5$  MeV

$\chi_{b0}(2P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\gamma\ T(2S)$	(4.6 ± 2.1) %		207
$\gamma\ T(1S)$	(9 ± 6) × 10 <sup>-3</sup>		743
$D^0 X$	< 8.2 %	90%	—
$\pi^+\pi^-K^+K^-\pi^0$	< 3.4 × 10 <sup>-5</sup>	90%	5064
$2\pi^+\pi^-K^-K_S^0$	< 5 × 10 <sup>-5</sup>	90%	5063
$2\pi^+\pi^-K^-K_S^0 2\pi^0$	< 2.2 × 10 <sup>-4</sup>	90%	5036
$2\pi^+2\pi^-2\pi^0$	< 2.4 × 10 <sup>-4</sup>	90%	5092
$2\pi^+2\pi^-K^+K^-$	< 1.5 × 10 <sup>-4</sup>	90%	5050
$2\pi^+2\pi^-K^+K^-\pi^0$	< 2.2 × 10 <sup>-4</sup>	90%	5035
$2\pi^+2\pi^-K^+K^-2\pi^0$	< 1.1 × 10 <sup>-3</sup>	90%	5019
$3\pi^+2\pi^-K^-K_S^0\pi^0$	< 7 × 10 <sup>-4</sup>	90%	5018
$3\pi^+3\pi^-$	< 7 × 10 <sup>-5</sup>	90%	5091
$3\pi^+3\pi^-2\pi^0$	< 1.2 × 10 <sup>-3</sup>	90%	5070
$3\pi^+3\pi^-K^+K^-$	< 1.5 × 10 <sup>-4</sup>	90%	5017
$3\pi^+3\pi^-K^+K^-\pi^0$	< 7 × 10 <sup>-4</sup>	90%	4999
$4\pi^+4\pi^-$	< 1.7 × 10 <sup>-4</sup>	90%	5069
$4\pi^+4\pi^-2\pi^0$	< 6 × 10 <sup>-4</sup>	90%	5039

<div><math>\chi_{b1}(2P)</math> <sup>[hhbb]</sup></div>	$I^G(J^{PC}) = 0^+(1^{++})$ $J$ needs confirmation.		
Mass $m = 10255.46 \pm 0.22 \pm 0.50$ MeV			
$m_{\chi_{b1}(2P)} - m_{\chi_{b0}(2P)} = 23.5 \pm 1.0$ MeV			
$\chi_{b1}(2P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor	$p$ ( MeV/c)
$\omega\ \mathcal{R}(1S)$	$(\ 1.63^{+0.40}_{-0.34})\%$		135
$\gamma\ \mathcal{R}(2S)$	$(19.9 \pm 1.9)\%$		230
$\gamma\ \mathcal{R}(1S)$	$(\ 9.2 \pm 0.8)\%$	1.1	764
$\pi\pi\chi_{b1}(1P)$	$(\ 9.1 \pm 1.3) \times 10^{-3}$		238
$D^0 X$	$(\ 8.8 \pm 1.7)\%$		—
$\pi^+ \pi^- K^+ K^- \pi^0$	$(\ 3.1 \pm 1.0) \times 10^{-4}$		5075
$2\pi^+ \pi^- K^- K_S^0$	$(\ 1.1 \pm 0.5) \times 10^{-4}$		5075
$2\pi^+ \pi^- K^- K_S^0 2\pi^0$	$(\ 7.7 \pm 3.2) \times 10^{-4}$		5047
$2\pi^+ 2\pi^- 2\pi^0$	$(\ 5.9 \pm 2.0) \times 10^{-4}$		5104
$2\pi^+ 2\pi^- K^+ K^-$	$(10 \pm 4) \times 10^{-5}$		5062
$2\pi^+ 2\pi^- K^+ K^- \pi^0$	$(\ 5.5 \pm 1.8) \times 10^{-4}$		5047
$2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	$(10 \pm 4) \times 10^{-4}$		5030
$3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	$(\ 6.7 \pm 2.6) \times 10^{-4}$		5029
$3\pi^+ 3\pi^-$	$(\ 1.2 \pm 0.4) \times 10^{-4}$		5103
$3\pi^+ 3\pi^- 2\pi^0$	$(\ 1.2 \pm 0.4) \times 10^{-3}$		5081
$3\pi^+ 3\pi^- K^+ K^-$	$(\ 2.0 \pm 0.8) \times 10^{-4}$		5029
$3\pi^+ 3\pi^- K^+ K^- \pi^0$	$(\ 6.1 \pm 2.2) \times 10^{-4}$		5011
$4\pi^+ 4\pi^-$	$(\ 1.7 \pm 0.6) \times 10^{-4}$		5080
$4\pi^+ 4\pi^- 2\pi^0$	$(\ 1.9 \pm 0.7) \times 10^{-3}$		5051

<b><math>\chi_{b2}(2P)</math></b> <sup>[hhbb]</sup>	$J^G(J^{PC}) = 0^+(2^{++})$ $J$ needs confirmation.
Mass $m = 10268.65 \pm 0.22 \pm 0.50$ MeV $m_{\chi_{b2}(2P)} - m_{\chi_{b1}(2P)} = 13.4 \pm 0.6$ MeV	

<b><math>\chi_{b2}(2P)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/Confidence level	$p$ (MeV/c)
$\omega\ T(1S)$	$(1.10^{+0.34}_{-0.30})\%$		194
$\gamma\ T(2S)$	$(10.6 \pm 2.6)\%$	S=2.0	242
$\gamma\ T(1S)$	$(7.0 \pm 0.7)\%$		777
$\pi\pi\chi_{b2}(1P)$	$(5.1 \pm 0.9) \times 10^{-3}$		229
$D^0 X$	$< 2.4\%$	CL=90%	–
$\pi^+\pi^-K^+K^-\pi^0$	$< 1.1 \times 10^{-4}$	CL=90%	5082
$2\pi^+\pi^-K^-K_S^0$	$< 9 \times 10^{-5}$	CL=90%	5082
$2\pi^+\pi^-K^-K_S^0 2\pi^0$	$< 7 \times 10^{-4}$	CL=90%	5054
$2\pi^+2\pi^-2\pi^0$	$(3.9 \pm 1.6) \times 10^{-4}$		5110
$2\pi^+2\pi^-K^+K^-$	$(9 \pm 4) \times 10^{-5}$		5068
$2\pi^+2\pi^-K^+K^-\pi^0$	$(2.4 \pm 1.1) \times 10^{-4}$		5054
$2\pi^+2\pi^-K^+K^-2\pi^0$	$(4.7 \pm 2.3) \times 10^{-4}$		5037
$3\pi^+2\pi^-K^-K_S^0\pi^0$	$< 4 \times 10^{-4}$	CL=90%	5036
$3\pi^+3\pi^-$	$(9 \pm 4) \times 10^{-5}$		5110
$3\pi^+3\pi^-2\pi^0$	$(1.2 \pm 0.4) \times 10^{-3}$		5088
$3\pi^+3\pi^-K^+K^-$	$(1.4 \pm 0.7) \times 10^{-4}$		5036
$3\pi^+3\pi^-K^+K^-\pi^0$	$(4.2 \pm 1.7) \times 10^{-4}$		5017
$4\pi^+4\pi^-$	$(9 \pm 5) \times 10^{-5}$		5087
$4\pi^+4\pi^-2\pi^0$	$(1.3 \pm 0.5) \times 10^{-3}$		5058

**$T(3S)$**

$I^G(J^{PC}) = 0^-(1^{--})$

Mass  $m = 10355.2 \pm 0.5$  MeV  
 $m_{T(3S)} - m_{T(2S)} = 331.50 \pm 0.13$  MeV  
 Full width  $\Gamma = 20.32 \pm 1.85$  keV  
 $\Gamma_{ee} = 0.443 \pm 0.008$  keV

<b><math>T(3S)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
$T(2S)$ anything	( 10.6 $\pm$ 0.8 ) %		296
$T(2S)\pi^+\pi^-$	( 2.82 $\pm$ 0.18 ) %	S=1.6	177
$T(2S)\pi^0\pi^0$	( 1.85 $\pm$ 0.14 ) %		190
$T(2S)\gamma\gamma$	( 5.0 $\pm$ 0.7 ) %		327
$T(2S)\pi^0$	< 5.1 $\times 10^{-4}$	CL=90%	298
$T(1S)\pi^+\pi^-$	( 4.37 $\pm$ 0.08 ) %		813
$T(1S)\pi^0\pi^0$	( 2.20 $\pm$ 0.13 ) %		816
$T(1S)\eta$	< 1 $\times 10^{-4}$	CL=90%	677
$T(1S)\pi^0$	< 7 $\times 10^{-5}$	CL=90%	846
$h_b(1P)\pi^0$	< 1.2 $\times 10^{-3}$	CL=90%	426
$h_b(1P)\pi^0 \rightarrow \gamma\eta_b(1S)\pi^0$	( 4.3 $\pm$ 1.4 ) $\times 10^{-4}$		—
$h_b(1P)\pi^+\pi^-$	< 1.2 $\times 10^{-4}$	CL=90%	353
$\tau^+\tau^-$	( 2.29 $\pm$ 0.30 ) %		4863
$\mu^+\mu^-$	( 2.18 $\pm$ 0.21 ) %	S=2.1	5177
$e^+e^-$	seen		5178
$ggg$	( 35.7 $\pm$ 2.6 ) %		—
$\gamma gg$	( 9.7 $\pm$ 1.8 ) $\times 10^{-3}$		—
$^2H$ anything	( 2.33 $\pm$ 0.33 ) $\times 10^{-5}$		—

<b>Radiative decays</b>			
$\gamma\chi_{b2}(2P)$	$(13.1 \pm 1.6)\%$	S=3.4	86
$\gamma\chi_{b1}(2P)$	$(12.6 \pm 1.2)\%$	S=2.4	99
$\gamma\chi_{b0}(2P)$	$(5.9 \pm 0.6)\%$	S=1.4	122
$\gamma\chi_{b2}(1P)$	$(9.9 \pm 1.3) \times 10^{-3}$	S=2.0	434
$\gamma A^0 \rightarrow \gamma$ hadrons	$< 8 \times 10^{-5}$	CL=90%	–
$\gamma\chi_{b1}(1P)$	$(9 \pm 5) \times 10^{-4}$	S=1.9	452
$\gamma\chi_{b0}(1P)$	$(2.7 \pm 0.4) \times 10^{-3}$		484
$\gamma\eta_b(2S)$	$< 6.2 \times 10^{-4}$	CL=90%	350
$\gamma\eta_b(1S)$	$(5.1 \pm 0.7) \times 10^{-4}$		913
$\gamma X \rightarrow \gamma + \geq 4$ prongs	[jjbb] $< 2.2 \times 10^{-4}$	CL=95%	–
$\gamma a_1^0 \rightarrow \gamma\mu^+\mu^-$	$< 5.5 \times 10^{-6}$	CL=90%	–
$\gamma a_1^0 \rightarrow \gamma\tau^+\tau^-$	[kkbb] $< 1.6 \times 10^{-4}$	CL=90%	–

<b>Lepton Family number (LF) violating modes</b>			
$e^\pm\tau^\mp$	LF $< 4.2 \times 10^{-6}$	CL=90%	5025
$\mu^\pm\tau^\mp$	LF $< 3.1 \times 10^{-6}$	CL=90%	5025

## Meson Summary Table

$\chi_{b1}(3P)$	$J^G(J^{PC}) = 0^+(1^+)$		
Mass $m = 10512.1 \pm 2.3$ MeV			
$\chi_{b1}(3P)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ ( MeV/c)	
$\Upsilon(1S)\gamma$	seen	999	
$\Upsilon(2S)\gamma$	seen	477	
$\Upsilon(3S)\gamma$	seen	156	

<b><math>\Upsilon(4S)</math></b> <b>or <math>\Upsilon(10580)</math></b>	$J^G(J^{PC}) = 0^-(1^{--})$
Mass $m = 10579.4 \pm 1.2$ MeV	
Full width $\Gamma = 20.5 \pm 2.5$ MeV	
$\Gamma_{ee} = 0.272 \pm 0.029$ keV    ( $S = 1.5$ )	

$\Upsilon(4S)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$B\bar{B}$	> 96 %	95%	326
$B^+B^-$	(51.4 $\pm$ 0.6 ) %		331
$D^+ \text{ anything } + \text{ c.c.}$	(17.8 $\pm$ 2.6 ) %		—
$B^0\bar{B}^0$	(48.6 $\pm$ 0.6 ) %		326
$J/\psi K_S^0 + (J/\psi, \eta_c) K_S^0$	< 4 $\times 10^{-7}$	90%	—
non- $B\bar{B}$	< 4 %	95%	—
$e^+e^-$	( 1.57 $\pm$ 0.08 ) $\times 10^{-5}$		5290
$\rho^+\rho^-$	< 5.7 $\times 10^{-6}$	90%	5233
$K^*(892)^0\bar{K}^0$	< 2.0 $\times 10^{-6}$	90%	5240
$J/\psi(1S)$ anything	< 1.9 $\times 10^{-4}$	95%	—
$D^{*+} \text{ anything } + \text{ c.c.}$	< 7.4 %	90%	5099
$\phi$ anything	( 7.1 $\pm$ 0.6 ) %		5240
$\phi\eta$	< 1.8 $\times 10^{-6}$	90%	5226
$\phi\eta'$	< 4.3 $\times 10^{-6}$	90%	5196
$\rho\eta$	< 1.3 $\times 10^{-6}$	90%	5247
$\rho\eta'$	< 2.5 $\times 10^{-6}$	90%	5217
$\Upsilon(1S)$ anything	< 4 $\times 10^{-3}$	90%	1053
$\Upsilon(1S)\pi^+\pi^-$	( 8.1 $\pm$ 0.6 ) $\times 10^{-5}$		1026
$\Upsilon(1S)\eta$	( 1.96 $\pm$ 0.28 ) $\times 10^{-4}$		924
$\Upsilon(2S)\pi^+\pi^-$	( 8.6 $\pm$ 1.3 ) $\times 10^{-5}$		468
$h_b(1P)\pi^+\pi^-$	not seen		600
$h_b(1P)\eta$	( 2.18 $\pm$ 0.21 ) $\times 10^{-3}$		390
$^2H$ anything	< 1.3 $\times 10^{-5}$	90%	—

<b><math>X(10610)^\pm</math></b>	$J^G(J^P) = 1^+(1^+)$
Mass $m = 10607.2 \pm 2.0$ MeV	
Full width $\Gamma = 18.4 \pm 2.4$ MeV	
$X(10610)^-$ decay modes are charge conjugates of the modes below.	

$X(10610)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Upsilon(1S)\pi^+$	seen	1077
$\Upsilon(2S)\pi^+$	seen	551
$\Upsilon(3S)\pi^+$	seen	207
$h_b(1P)\pi^+$	seen	671
$h_b(2P)\pi^+$	seen	313

<div><math>X(10610)^0</math></div>	$I^G(J^P) = 1^+(1^+)$		
Mass $m = 10609 \pm 6$ MeV			
$X(10610)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)	
$\Upsilon(1S)\pi^0$	not seen	1079	
$\Upsilon(2S)\pi^0$	seen	554	
$\Upsilon(3S)\pi^0$	seen	212	

<b><math>\Upsilon(10860)</math></b>	$J^G(J^{PC}) = 0^-(1^{--})$
Mass $m = 10891 \pm 4$ MeV	
Full width $\Gamma = 54 \pm 7$ MeV	
$\Gamma_{ee} = 0.31 \pm 0.07$ keV (S = 1.3)	

$\Upsilon(10860)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$B\bar{B}X$	( 76.2 $\pm$ 2.7 $\pm$ 4.0 ) %		—
$B\bar{B}$	( 5.5 $\pm$ 1.0 ) %		1334
$B\bar{B}^* + \text{c.c.}$	( 13.7 $\pm$ 1.6 ) %		—
$B_s^*\bar{B}^*$	( 38.1 $\pm$ 3.4 ) %		1141
$B\bar{B}^{(*)}\pi$	< 19.7 %	90%	1031
$B\bar{B}\pi$	( 0.0 $\pm$ 1.2 ) %		1031
$B^*\bar{B}\pi + B\bar{B}^*\pi$	( 7.3 $\pm$ 2.3 ) %		—
$B^*\bar{B}^*\pi$	( 1.0 $\pm$ 1.4 ) %		761
$B\bar{B}\pi\pi$	< 8.9 %	90%	580
$B_s^{(*)}\bar{B}_s^{(*)}$	( 20.1 $\pm$ 3.1 ) %		923
$B_s\bar{B}_s$	( 5 $\pm$ 5 ) $\times 10^{-3}$		923
$B_s\bar{B}_s^* + \text{c.c.}$	( 1.35 $\pm$ 0.32 ) %		—
$B_s^*\bar{B}_s^*$	( 17.6 $\pm$ 2.7 ) %		572
no open-bottom	( 3.8 $\pm$ 5.0 $\pm$ 0.5 ) %		—
$e^+e^-$	( 5.7 $\pm$ 1.5 ) $\times 10^{-6}$		5446
$K^*(892)^0\bar{K}^0$	< 1.0 $\times 10^{-5}$	90%	5398
$\Upsilon(1S)\pi^+\pi^-$	( 5.3 $\pm$ 0.6 ) $\times 10^{-3}$		1311
$\Upsilon(2S)\pi^+\pi^-$	( 7.8 $\pm$ 1.3 ) $\times 10^{-3}$		789
$\Upsilon(3S)\pi^+\pi^-$	( 4.8 $\pm$ 1.9 $\pm$ 1.7 ) $\times 10^{-3}$		446
$\Upsilon(1S)K^+K^-$	( 6.1 $\pm$ 1.8 ) $\times 10^{-4}$		966
$h_b(1P)\pi^+\pi^-$	( 3.5 $\pm$ 1.0 $\pm$ 1.3 ) $\times 10^{-3}$		908
$h_b(2P)\pi^+\pi^-$	( 6.0 $\pm$ 2.1 $\pm$ 1.8 ) $\times 10^{-3}$		550
$\chi_{b0}(1P)\pi^+\pi^-\pi^0$	< 6.3 $\times 10^{-3}$	90%	900
$\chi_{b0}(1P)\omega$	< 3.9 $\times 10^{-3}$	90%	640
$\chi_{b0}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega}$	< 4.8 $\times 10^{-3}$	90%	—
$\chi_{b1}(1P)\pi^+\pi^-\pi^0$	( 1.85 $\pm$ 0.33 ) $\times 10^{-3}$		867
$\chi_{b1}(1P)\omega$	( 1.57 $\pm$ 0.30 ) $\times 10^{-3}$		591
$\chi_{b1}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega}$	( 5.2 $\pm$ 1.9 ) $\times 10^{-4}$		—
$\chi_{b2}(1P)\pi^+\pi^-\pi^0$	( 1.17 $\pm$ 0.30 ) $\times 10^{-3}$		847
$\chi_{b2}(1P)\omega$	( 6.0 $\pm$ 2.7 ) $\times 10^{-4}$		561
$\chi_{b2}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega}$	( 6 $\pm$ 4 ) $\times 10^{-4}$		—
$\gamma X_b \rightarrow \gamma \Upsilon(1S)\omega$	< 3.8 $\times 10^{-5}$	90%	—

## Inclusive Decays.

These decay modes are submodes of one or more of the decay modes above.

$\phi$ anything	( 13.8 $\pm$ 2.4 $\pm$ 1.7 ) %	—
$D^0$ anything + c.c.	(108 $\pm$ 8 ) %	—
$D_s$ anything + c.c.	( 46 $\pm$ 6 ) %	—
$J/\psi$ anything	( 2.06 $\pm$ 0.21 ) %	—
$B^0$ anything + c.c.	( 77 $\pm$ 8 ) %	—
$B^+$ anything + c.c.	( 72 $\pm$ 6 ) %	—

<div><div><math>\Upsilon(11020)</math></div></div>	$I^G(J^{PC}) = 0^-(1^--)$	
Mass $m = 10987.5^{+11.0}_{-3.4}$ MeV		
Full width $\Gamma = 61^{+9}_{-28}$ MeV		
$\Gamma_{ee} = 0.130 \pm 0.030$ keV		
<b><math>\Upsilon(11020)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$e^+e^-$	$(2.1^{+1.1}_{-0.6}) \times 10^{-6}$	5494

# Meson Summary Table

## NOTES

In this Summary Table:

When a quantity has “(S = ...)” to its right, the error on the quantity has been enlarged by the “scale factor” S, defined as  $S = \sqrt{\chi^2/(N-1)}$ , where N is the number of measurements used in calculating the quantity. We do this when  $S > 1$ , which often indicates that the measurements are inconsistent. When  $S > 1.25$ , we also show in the Particle Listings an ideogram of the measurements. For more about S, see the Introduction.

A decay momentum  $p$  is given for each decay mode. For a 2-body decay,  $p$  is the momentum of each decay product in the rest frame of the decaying particle. For a 3-or-more-body decay,  $p$  is the largest momentum any of the products can have in this frame.

- [a] See the “Note on  $\pi^\pm \rightarrow \ell^\pm \nu \gamma$  and  $K^\pm \rightarrow \ell^\pm \nu \gamma$  Form Factors” in the  $\pi^\pm$  Particle Listings for definitions and details.
- [b] Measurements of  $\Gamma(e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$  always include decays with  $\gamma$ ’s, and measurements of  $\Gamma(e^+ \nu_e \gamma)$  and  $\Gamma(\mu^+ \nu_\mu \gamma)$  never include low-energy  $\gamma$ ’s. Therefore, since no clean separation is possible, we consider the modes with  $\gamma$ ’s to be subreactions of the modes without them, and let  $[\Gamma(e^+ \nu_e) + \Gamma(\mu^+ \nu_\mu)]/\Gamma_{\text{total}} = 100\%$ .
- [c] See the  $\pi^\pm$  Particle Listings for the energy limits used in this measurement; low-energy  $\gamma$ ’s are not included.
- [d] Derived from an analysis of neutrino-oscillation experiments.
- [e] Astrophysical and cosmological arguments give limits of order  $10^{-13}$ ; see the  $\pi^0$  Particle Listings.
- [f] C parity forbids this to occur as a single-photon process.
- [g] See the “Note on scalar mesons” in the  $f_0(500)$  Particle Listings. The interpretation of this entry as a particle is controversial.
- [h] See the “Note on  $\rho(770)$ ” in the  $\rho(770)$  Particle Listings.
- [i] The  $\omega\rho$  interference is then due to  $\omega\rho$  mixing only, and is expected to be small. If  $e\mu$  universality holds,  $\Gamma(\rho^0 \rightarrow \mu^+ \mu^-) = \Gamma(\rho^0 \rightarrow e^+ e^-) \times 0.99785$ .
- [j] See the “Note on scalar mesons” in the  $f_0(500)$  Particle Listings.
- [k] See the “Note on  $a_1(1260)$ ” in the  $a_1(1260)$  Particle Listings in PDG 06, Journal of Physics **G33** 1 (2006).
- [l] This is only an educated guess; the error given is larger than the error on the average of the published values. See the Particle Listings for details.
- [n] See the “Note on non- $q\bar{q}$  mesons” in the Particle Listings in PDG 06, Journal of Physics **G33** 1 (2006).
- [o] See the “Note on the  $\eta(1405)$ ” in the  $\eta(1405)$  Particle Listings.
- [p] See the “Note on the  $f_1(1420)$ ” in the  $\eta(1405)$  Particle Listings.
- [q] See also the  $\omega(1650)$  Particle Listings.
- [r] See the “Note on the  $\rho(1450)$  and the  $\rho(1700)$ ” in the  $\rho(1700)$  Particle Listings.
- [s] See also the  $\omega(1420)$  Particle Listings.
- [t] See the “Note on  $f_0(1710)$ ” in the  $f_0(1710)$  Particle Listings in 2004 edition of *Review of Particle Physics*.
- [u] See the note in the  $K^\pm$  Particle Listings.
- [v] Neglecting photon channels. See, e.g., A. Pais and S.B. Treiman, Phys. Rev. **D12**, 2744 (1975).
- [x] The definition of the slope parameters of the  $K \rightarrow 3\pi$  Dalitz plot is as follows (see also “Note on Dalitz Plot Parameters for  $K \rightarrow 3\pi$  Decays” in the  $K^\pm$  Particle Listings):
 
$$|M|^2 = 1 + g(s_3 - s_0)/m_{\pi^+}^2 + \dots$$
- [y] For more details and definitions of parameters see the Particle Listings.
- [z] See the  $K^\pm$  Particle Listings for the energy limits used in this measurement.
- [aa] Most of this radiative mode, the low-momentum  $\gamma$  part, is also included in the parent mode listed without  $\gamma$ ’s.
- [bb] Structure-dependent part.
- [cc] Direct-emission branching fraction.
- [dd] Violates angular-momentum conservation.
- [ee] Derived from measured values of  $\phi_{+-}$ ,  $\phi_{00}$ ,  $|\eta|$ ,  $|m_{K_L^0} - m_{K_S^0}|$ , and  $\tau_{K_S^0}$ , as described in the introduction to “Tests of Conservation Laws.”
- [ff] The CP-violation parameters are defined as follows (see also “Note on CP Violation in  $K_S \rightarrow 3\pi$ ” and “Note on CP Violation in  $K_L^0$  Decay” in the Particle Listings):

$$\eta_{+-} = |\eta_{+-}|e^{i\phi_{+-}} = \frac{A(K_L^0 \rightarrow \pi^+ \pi^-)}{A(K_S^0 \rightarrow \pi^+ \pi^-)} = \epsilon + \epsilon'$$

$$\eta_{00} = |\eta_{00}|e^{i\phi_{00}} = \frac{A(K_L^0 \rightarrow \pi^0 \pi^0)}{A(K_S^0 \rightarrow \pi^0 \pi^0)} = \epsilon - 2\epsilon'$$

$$\delta = \frac{\Gamma(K_L^0 \rightarrow \pi^- \ell^+ \nu) - \Gamma(K_L^0 \rightarrow \pi^+ \ell^- \nu)}{\Gamma(K_L^0 \rightarrow \pi^- \ell^+ \nu) + \Gamma(K_L^0 \rightarrow \pi^+ \ell^- \nu)},$$

$$\text{Im}(\eta_{+-0})^2 = \frac{\Gamma(K_S^0 \rightarrow \pi^+ \pi^- \pi^0)_{CP \text{ viol.}}}{\Gamma(K_L^0 \rightarrow \pi^+ \pi^- \pi^0)},$$

$$\text{Im}(\eta_{000})^2 = \frac{\Gamma(K_S^0 \rightarrow \pi^0 \pi^0 \pi^0)}{\Gamma(K_L^0 \rightarrow \pi^0 \pi^0 \pi^0)}.$$

where for the last two relations CPT is assumed valid, i.e.,  $\text{Re}(\eta_{+-0}) \simeq 0$  and  $\text{Re}(\eta_{000}) \simeq 0$ .

- [gg] See the  $K_S^0$  Particle Listings for the energy limits used in this measurement.
- [hh] The value is for the sum of the charge states or particle/antiparticle states indicated.
- [ii]  $\text{Re}(\epsilon'/\epsilon) = \epsilon'/\epsilon$  to a very good approximation provided the phases satisfy CPT invariance.
- [jj] This mode includes gammas from inner bremsstrahlung but not the direct emission mode  $K_L^0 \rightarrow \pi^+ \pi^- \gamma$  (DE).
- [kk] See the  $K_L^0$  Particle Listings for the energy limits used in this measurement.
- [ll] Allowed by higher-order electroweak interactions.
- [nn] Violates CP in leading order. Test of direct CP violation since the indirect CP-violating and CP-conserving contributions are expected to be suppressed.
- [oo] See the “Note on  $f_0(1370)$ ” in the  $f_0(1370)$  Particle Listings and in the 1994 edition.
- [pp] See the note in the  $L(1770)$  Particle Listings in Reviews of Modern Physics **56** S1 (1984), p. S200. See also the “Note on  $K_2(1770)$  and the  $K_2(1820)$ ” in the  $K_2(1770)$  Particle Listings.
- [qq] See the “Note on  $K_2(1770)$  and the  $K_2(1820)$ ” in the  $K_2(1770)$  Particle Listings.
- [rr] This result applies to  $Z^0 \rightarrow c\bar{c}$  decays only. Here  $\ell^+$  is an average (not a sum) of  $e^+$  and  $\mu^+$  decays.
- [ss] See the Particle Listings for the (complicated) definition of this quantity.
- [tt] The branching fraction for this mode may differ from the sum of the submodes that contribute to it, due to interference effects. See the relevant papers in the Particle Listings.
- [uu] These subfractions of the  $K^- 2\pi^+$  mode are uncertain: see the Particle Listings.
- [vv] Submodes of the  $D^+ \rightarrow K^- 2\pi^+ \pi^0$  and  $K_S^0 2\pi^+ \pi^-$  modes were studied by ANJOS 92C and COFFMAN 92B, but with at most 142 events for the first mode and 229 for the second – not enough for precise results. With nothing new for 18 years, we refer to our 2008 edition, Physics Letters **B667** 1 (2008), for those results.
- [xx] The unseen decay modes of the resonances are included.
- [yy] This is *not* a test for the  $\Delta C=1$  weak neutral current, but leads to the  $\pi^+ \ell^+ \ell^-$  final state.
- [zz] This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.
- [aaa] In the 2010 *Review*, the values for these quantities were given using a measure of the asymmetry that was inconsistent with the usual definition.
- [bbb] This value is obtained by subtracting the branching fractions for 2-, 4- and 6-prongs from unity.
- [ccc] This is the sum of our  $K^- 2\pi^+ \pi^-$ ,  $K^- 2\pi^+ \pi^- \pi^0$ ,  $K^0 2\pi^+ 2\pi^-$ ,  $K^+ 2K^- \pi^+$ ,  $2\pi^+ 2\pi^-$ ,  $2\pi^+ 2\pi^- \pi^0$ ,  $K^+ K^- \pi^+ \pi^-$ , and  $K^+ K^- \pi^+ \pi^- \pi^0$ , branching fractions.
- [ddd] This is the sum of our  $K^- 3\pi^+ 2\pi^-$  and  $3\pi^+ 3\pi^-$  branching fractions.
- [eee] The branching fractions for the  $K^- e^+ \nu_e$ ,  $K^*(892)^- e^+ \nu_e$ ,  $\pi^- e^+ \nu_e$ , and  $\rho^- e^+ \nu_e$  modes add up to  $6.19 \pm 0.17\%$ .
- [fff] This is a doubly Cabibbo-suppressed mode.
- [ggg] The two experiments measuring this fraction are in serious disagreement. See the Particle Listings.
- [hhh] Submodes of the  $D^0 \rightarrow K_S^0 \pi^+ \pi^- \pi^0$  mode with a  $K^*$  and/or  $\rho$  were studied by COFFMAN 92B, but with only 140 events. With nothing new

# Meson Summary Table

for 18 years, we refer to our 2008 edition, Physics Letters <b>B667</b> 1 (2008), for those results.	
[iii]	This branching fraction includes all the decay modes of the resonance in the final state.
[jjj]	This limit is for either $D^0$ or $\overline{D}^0$ to $p e^-$ .
[kkk]	This limit is for either $D^0$ or $\overline{D}^0$ to $\overline{p} e^+$ .
[lll]	This is the purely $e^+$ semileptonic branching fraction: the $e^+$ fraction from $\tau^+$ decays has been subtracted off. The sum of our (non- $\tau$ ) $e^+$ exclusive fractions — an $e^+ \nu_e$ with an $\eta$ , $\eta'$ , $\phi$ , $K^0$ , $K^{*0}$ , or $f_0(980)$ — is $7.0 \pm 0.4\%$
[nnn]	This fraction includes $\eta$ from $\eta'$ decays.
[ooo]	Two times (to include $\mu$ decays) the $\eta' e^+ \nu_e$ branching fraction, plus the $\eta' \pi^+$ , $\eta' \rho^+$ , and $\eta' K^+$ fractions, is $(18.6 \pm 2.3)\%$ , which considerably exceeds the inclusive $\eta'$ fraction of $(11.7 \pm 1.8)\%$ . Our best guess is that the $\eta' \rho^+$ fraction, $(12.5 \pm 2.2)\%$ , is too large.
[ppp]	This branching fraction includes all the decay modes of the final-state resonance.
[qqq]	A test for $u\overline{u}$ or $d\overline{d}$ content in the $D_s^+$ . Neither Cabibbo-favored nor Cabibbo-suppressed decays can contribute, and $\omega$ – $\phi$ mixing is an unlikely explanation for any fraction above about $2 \times 10^{-4}$ .
[rrr]	We decouple the $D_s^+ \rightarrow \phi \pi^+$ branching fraction obtained from mass projections (and used to get some of the other branching fractions) from the $D_s^+ \rightarrow \phi \pi^+$ , $\phi \rightarrow K^+ K^-$ branching fraction obtained from the Dalitz-plot analysis of $D_s^+ \rightarrow K^+ K^- \pi^+$ . That is, the ratio of these two branching fractions is not exactly the $\phi \rightarrow K^+ K^-$ branching fraction 0.491.
[sss]	This is the average of a model-independent and a $K$ -matrix parametrization of the $\pi^+ \pi^-$ $S$ -wave and is a sum over several $f_0$ mesons.
[ttt]	An $\ell$ indicates an $e$ or a $\mu$ mode, not a sum over these modes.
[uuu]	An $CP(\pm 1)$ indicates the $CP=+1$ and $CP=-1$ eigenstates of the $D^0$ – $\overline{D}^0$ system.
[vvv]	$D$ denotes $D^0$ or $\overline{D}^0$ .
[xxx]	$D_{CP+}^{*0}$ decays into $D^0 \pi^0$ with the $D^0$ reconstructed in $CP$ -even eigenstates $K^+ K^-$ and $\pi^+ \pi^-$ .
[yyy]	$\overline{D}^{**}$ represents an excited state with mass $2.2 < M < 2.8$ GeV/ $c^2$ .
[zzz]	$X(3872)^+$ is a hypothetical charged partner of the $X(3872)$ .
[aaa]	$\Theta(1710)^{++}$ is a possible narrow pentaquark state and $G(2220)$ is a possible glueball resonance.
[baa]	$(\overline{A}_c^- p)_s$ denotes a low-mass enhancement near 3.35 GeV/ $c^2$ .
[caa]	Stands for the possible candidates of $K^*(1410)$ , $K_0^*(1430)$ and $K_2^*(1430)$ .
[daa]	$B^0$ and $B_s^0$ contributions not separated. Limit is on weighted average of the two decay rates.
[eaa]	This decay refers to the coherent sum of resonant and nonresonant $J^P = 0^+ K \pi$ components with $1.60 < m_{K\pi} < 2.15$ GeV/ $c^2$ .
[faa]	$X(214)$ is a hypothetical particle of mass 214 MeV/ $c^2$ reported by the HyperCP experiment, Physical Review Letters <b>94</b> 021801 (2005)
[gga]	$\Theta(1540)^+$ denotes a possible narrow pentaquark state.
[hha]	Here $S$ and $P$ are the hypothetical scalar and pseudoscalar particles with masses of 2.5 GeV/ $c^2$ and 214.3 MeV/ $c^2$ , respectively.
[jia]	These values are model dependent.
[jja]	Here “anything” means at least one particle observed.
[kaa]	This is a $B(B^0 \rightarrow D^{*-} \ell^+ \nu_\ell)$ value.
[laa]	$D^{**}$ stands for the sum of the $D(1^1 P_1)$ , $D(1^3 P_0)$ , $D(1^3 P_1)$ , $D(1^3 P_2)$ , $D(2^1 S_0)$ , and $D(2^1 S_1)$ resonances.
[naa]	$D^{(*)} \overline{D}^{(*)}$ stands for the sum of $D^* \overline{D}^*$ , $D^* \overline{D}$ , $D \overline{D}^*$ , and $D \overline{D}$ .
[oaa]	$X(3915)$ denotes a near-threshold enhancement in the $\omega J/\psi$ mass spectrum.
[ppa]	Inclusive branching fractions have a multiplicity definition and can be greater than 100%.
[qqa]	$D_j$ represents an unresolved mixture of pseudoscalar and tensor $D^{**}$ ( $P$ -wave) states.
[rra]	Not a pure measurement. See note at head of $B_s^0$ Decay Modes.
[saa]	For $E_\gamma > 100$ MeV.
[taa]	Includes $p \overline{p} \pi^+ \pi^- \gamma$ and excludes $p \overline{p} \eta$ , $p \overline{p} \omega$ , $p \overline{p} \eta'$ .
[uaa]	For a narrow state $A$ with mass less than 960 MeV.
[vaa]	For a narrow scalar or pseudoscalar $A^0$ with mass 0.21–3.0 GeV.
[xaa]	For a narrow resonance in the range $2.2 < M(X) < 2.8$ GeV.
[yaa]	$J^{PC}$ known by production in $e^+ e^-$ via single photon annihilation. $I^G$ is not known; interpretation of this state as a single resonance is unclear because of the expectation of substantial threshold effects in this energy region.
[zza]	$2m_\tau < M(\tau^+ \tau^-) < 9.2$ GeV
[aab]	$2 \text{ GeV} < m_{K^+ K^-} < 3$ GeV
[bbb]	$X = \text{scalar with } m < 8.0$ GeV
[ccb]	$X \overline{X} = \text{vectors with } m < 3.1$ GeV
[ddb]	$X$ and $\overline{X} = \text{zero spin with } m < 4.5$ GeV
[eeb]	$1.5 \text{ GeV} < m_X < 5.0$ GeV
[ffb]	$201 \text{ MeV} < M(\mu^+ \mu^-) < 3565$ MeV
[ggb]	$0.5 \text{ GeV} < m_X < 9.0$ GeV, where $m_X$ is the invariant mass of the hadronic final state.
[hbb]	Spectroscopic labeling for these states is theoretical, pending experimental information.
[iib]	$1.5 \text{ GeV} < m_X < 5.0$ GeV
[jib]	$1.5 \text{ GeV} < m_X < 5.0$ GeV
[kbb]	For $m_{\tau^+ \tau^-}$ in the ranges 4.03–9.52 and 9.61–10.10 GeV.

- Indicates particles that appear in the preceding Meson Summary Table. We do not regard the other entries as being established.

LIGHT UNFLAVORED ( $S = C = B = 0$ )				STRANGE ( $S = \pm 1, C = B = 0$ )		CHARMED, STRANGE ( $C = S = \pm 1$ )		$c\bar{c}$ $I^G(J^{PC})$	
$I^G(J^{PC})$		$I^G(J^{PC})$		$I(J^P)$		$I(J^P)$			
<ul style="list-style-type: none"><li><math>\pi^\pm</math></li><li><math>\pi^0</math></li><li><math>\eta</math></li><li><math>f_0(500)</math></li><li><math>\rho(770)</math></li><li><math>\omega(782)</math></li><li><math>\eta'(958)</math></li><li><math>f_0(980)</math></li><li><math>a_0(980)</math></li><li><math>\phi(1020)</math></li><li><math>h_1(1170)</math></li><li><math>b_1(1235)</math></li><li><math>a_1(1260)</math></li><li><math>f_2(1270)</math></li><li><math>f_1(1285)</math></li><li><math>\eta(1295)</math></li><li><math>\pi(1300)</math></li><li><math>a_2(1320)</math></li><li><math>f_0(1370)</math></li><li><math>h_1(1380)</math></li><li><math>\pi_1(1400)</math></li><li><math>\eta(1405)</math></li><li><math>f_1(1420)</math></li><li><math>\omega(1420)</math></li><li><math>f_2(1430)</math></li><li><math>a_0(1450)</math></li><li><math>\rho(1450)</math></li><li><math>\eta(1475)</math></li><li><math>f_0(1500)</math></li><li><math>f_1(1510)</math></li><li><math>f_2'(1525)</math></li><li><math>f_2(1565)</math></li><li><math>\rho(1570)</math></li><li><math>h_1(1595)</math></li><li><math>\pi_1(1600)</math></li><li><math>a_1(1640)</math></li><li><math>f_2(1640)</math></li><li><math>\eta_2(1645)</math></li><li><math>\omega(1650)</math></li><li><math>\omega_3(1670)</math></li><li><math>\pi_2(1670)</math></li><li><math>\phi(1680)</math></li></ul>	<ul style="list-style-type: none"><li><math>\rho_3(1690)</math></li><li><math>\rho(1700)</math></li><li><math>a_2(1700)</math></li><li><math>f_0(1710)</math></li><li><math>\eta(1760)</math></li><li><math>f_2(1810)</math></li><li><math>X(1835)</math></li><li><math>X(1840)</math></li><li><math>a_1(1420)</math></li><li><math>\phi_3(1850)</math></li><li><math>\eta_2(1870)</math></li><li><math>\pi_2(1880)</math></li><li><math>\rho(1900)</math></li><li><math>f_2(1910)</math></li><li><math>a_0(1950)</math></li><li><math>f_2(1950)</math></li><li><math>\rho_3(1990)</math></li><li><math>f_2(2010)</math></li><li><math>f_0(2020)</math></li><li><math>a_4(2040)</math></li><li><math>f_4(2050)</math></li><li><math>\pi_2(2100)</math></li><li><math>f_0(2100)</math></li><li><math>f_2(2150)</math></li><li><math>\rho(2150)</math></li><li><math>\phi(2170)</math></li><li><math>f_0(2200)</math></li><li><math>f_J(2220)</math></li><li><math>\eta(2225)</math></li><li><math>\rho_3(2250)</math></li><li><math>f_2(2300)</math></li><li><math>f_4(2300)</math></li><li><math>f_0(2330)</math></li><li><math>f_2(2340)</math></li><li><math>\rho_5(2350)</math></li><li><math>a_6(2450)</math></li><li><math>f_6(2510)</math></li></ul>	<ul style="list-style-type: none"><li><math>1^-(0^-)</math></li><li><math>1^-(0^-+)</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(0^{++})</math></li><li><math>1^+(1^{--})</math></li><li><math>0^-(1^{--})</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(0^{++})</math></li><li><math>1^-(0^{++})</math></li><li><math>0^-(1^{+-})</math></li><li><math>1^+(1^{+-})</math></li><li><math>1^-(1^{++})</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(1^{++})</math></li><li><math>0^-(1^{--})</math></li><li><math>0^-(0^-+)</math></li><li><math>1^-(2^{++})</math></li><li><math>0^+(0^{++})</math></li><li><math>h_1(1380)</math></li><li><math>1^-(1^{+-})</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(1^{++})</math></li><li><math>0^-(1^{--})</math></li><li><math>0^+(2^{++})</math></li><li><math>1^-(0^{++})</math></li><li><math>1^-(1^{+-})</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(1^{++})</math></li><li><math>0^-(1^{--})</math></li><li><math>0^+(2^{++})</math></li><li><math>1^-(0^{++})</math></li><li><math>1^-(1^{+-})</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(2^{++})</math></li><li><math>0^-(1^{--})</math></li><li><math>0^-(3^{--})</math></li><li><math>1^-(2^-+)</math></li><li><math>0^-(1^-+)</math></li></ul>	<ul style="list-style-type: none"><li><math>1^+(3^{--})</math></li><li><math>1^+(1^{--})</math></li><li><math>1^-(2^{++})</math></li><li><math>0^+(0^-+)</math></li><li><math>0^+(0^-+)</math></li><li><math>1^-(0^-+)</math></li><li><math>0^+(2^{++})</math></li><li><math>?^?(0^-+)</math></li><li><math>?^?(?^{??})</math></li><li><math>1^-(1^{++})</math></li><li><math>0^-(3^{--})</math></li><li><math>0^+(2^-+)</math></li><li><math>1^-(2^-+)</math></li><li><math>1^+(1^-+)</math></li><li><math>0^+(2^{++})</math></li><li><math>1^-(0^{++})</math></li><li><math>0^+(2^{++})</math></li><li><math>1^-(4^{++})</math></li><li><math>0^+(4^{++})</math></li><li><math>1^-(2^-+)</math></li><li><math>0^+(0^{++})</math></li><li><math>0^+(2^{++})</math></li><li><math>1^-(2^-+)</math></li><li><math>0^+(0^{++})</math></li><li><math>0^+(2^{++})</math></li><li><math>1^-(6^{++})</math></li><li><math>0^+(6^{++})</math></li></ul>	<ul style="list-style-type: none"><li><math>K^\pm</math></li><li><math>K^0</math></li><li><math>K_S^0</math></li><li><math>K_L^0</math></li><li><math>K_0^*(800)</math></li><li><math>K^*(892)</math></li><li><math>K_1(1270)</math></li><li><math>K_1(1400)</math></li><li><math>K^*(1410)</math></li><li><math>K_0^*(1430)</math></li><li><math>K_2^*(1430)</math></li><li><math>K(1460)</math></li><li><math>K_2(1580)</math></li><li><math>K(1630)</math></li><li><math>K_1(1650)</math></li><li><math>K^*(1680)</math></li><li><math>K_2(1770)</math></li><li><math>K_3^*(1780)</math></li><li><math>K_2(1820)</math></li><li><math>K(1830)</math></li><li><math>K_0^*(1950)</math></li><li><math>K_2^*(1980)</math></li><li><math>K_4^*(2045)</math></li><li><math>K_2(2250)</math></li><li><math>K_3(2320)</math></li><li><math>K_5^*(2380)</math></li><li><math>K_4(2500)</math></li><li><math>K(3100)</math></li></ul>	<ul style="list-style-type: none"><li><math>1/2(0^-)</math></li><li><math>1/2(0^-)</math></li><li><math>1/2(0^-)</math></li><li><math>1/2(0^-)</math></li><li><math>1/2(0^+)</math></li><li><math>1/2(1^-)</math></li><li><math>1/2(1^+)</math></li><li><math>1/2(1^+)</math></li><li><math>1/2(1^-)</math></li><li><math>1/2(0^+)</math></li><li><math>1/2(2^+)</math></li><li><math>1/2(0^-)</math></li><li><math>1/2(2^-)</math></li><li><math>1/2(2^-)</math></li><li><math>1/2(2^-)</math></li><li><math>1/2(0^-)</math></li><li><math>1/2(0^+)</math></li><li><math>1/2(2^+)</math></li><li><math>1/2(4^+)</math></li><li><math>1/2(2^-)</math></li><li><math>1/2(3^+)</math></li><li><math>1/2(5^-)</math></li><li><math>1/2(4^-)</math></li><li><math>?^?(?^{??})</math></li></ul>	<ul style="list-style-type: none"><li><math>D_s^\pm</math></li><li><math>D_s^{*\pm}</math></li><li><math>D_{s0}^*(2317)^\pm</math></li><li><math>D_{s1}(2460)^\pm</math></li><li><math>D_{s1}(2536)^\pm</math></li><li><math>D_{s2}(2573)</math></li><li><math>D_{s1}^*(2700)^\pm</math></li><li><math>D_{s1}^*(2860)^\pm</math></li><li><math>D_{s1}^*(2860)^\pm</math></li><li><math>D_{sJ}(3040)^\pm</math></li></ul>	<ul style="list-style-type: none"><li><math>0(0^-)</math></li><li><math>0(?^?)</math></li><li><math>0(0^+)</math></li><li><math>0(1^+)</math></li><li><math>0(1^+)</math></li><li><math>0(2^+)</math></li><li><math>0(1^-)</math></li><li><math>0(1^-)</math></li><li><math>0(3^-)</math></li><li><math>0(?^?)</math></li></ul>	<ul style="list-style-type: none"><li><math>\eta_c(1S)</math></li><li><math>J/\psi(1S)</math></li><li><math>\chi_{c0}(1P)</math></li><li><math>\chi_{c1}(1P)</math></li><li><math>h_c(1P)</math></li><li><math>\chi_{c2}(1P)</math></li><li><math>\eta_c(2S)</math></li><li><math>\psi(2S)</math></li><li><math>\psi(3770)</math></li><li><math>\psi(3823)</math></li><li><math>X(3872)</math></li><li><math>X(3900)</math></li><li><math>X(3915)</math></li><li><math>\chi_{c2}(2P)</math></li><li><math>X(3940)</math></li><li><math>X(4020)</math></li><li><math>\psi(4040)</math></li><li><math>X(4050)^\pm</math></li><li><math>X(4055)^\pm</math></li><li><math>X(4140)</math></li></ul>	
				CHARMED ( $C = \pm 1$ )		BOTTOM, STRANGE ( $B = \pm 1, S = \mp 1$ )		$b\bar{b}$	
				<ul style="list-style-type: none"><li><math>D^\pm</math></li><li><math>D^0</math></li><li><math>D^*(2007)^0</math></li><li><math>D^*(2010)^\pm</math></li><li><math>D_0^*(2400)^0</math></li><li><math>D_0^*(2400)^\pm</math></li><li><math>D_1(2420)^0</math></li><li><math>D_1(2420)^\pm</math></li><li><math>D_1(2430)^0</math></li><li><math>D_2^*(2460)^0</math></li><li><math>D_2^*(2460)^\pm</math></li><li><math>D(2550)^0</math></li><li><math>D_J(2600)</math></li><li><math>D^*(2640)^\pm</math></li><li><math>D(2740)^0</math></li><li><math>D(2750)</math></li><li><math>D(3000)^0</math></li></ul>	<ul style="list-style-type: none"><li><math>1/2(0^-)</math></li><li><math>1/2(0^-)</math></li><li><math>1/2(1^-)</math></li><li><math>1/2(1^-)</math></li><li><math>1/2(0^+)</math></li><li><math>1/2(0^+)</math></li><li><math>1/2(1^+)</math></li><li><math>1/2(?^?)</math></li><li><math>1/2(1^+)</math></li><li><math>1/2(2^+)</math></li><li><math>1/2(2^+)</math></li><li><math>1/2(?^?)</math></li><li><math>1/2(?^?)</math></li><li><math>1/2(?^?)</math></li><li><math>1/2(3^-)</math></li><li><math>1/2(?^?)</math></li></ul>	<ul style="list-style-type: none"><li><math>B_s^0</math></li><li><math>B_s^+</math></li><li><math>B_{s1}(5830)^0</math></li><li><math>B_{s2}^*(5840)^0</math></li><li><math>B_{sJ}^*(5850)</math></li></ul>	<ul style="list-style-type: none"><li><math>0(0^-)</math></li><li><math>0(1^-)</math></li><li><math>0(1^+)</math></li><li><math>0(2^+)</math></li><li><math>?(?^?)</math></li></ul>	<ul style="list-style-type: none"><li><math>\eta_b(1S)</math></li><li><math>\Upsilon(1S)</math></li><li><math>\chi_{b0}(1P)</math></li><li><math>\chi_{b1}(1P)</math></li><li><math>h_b(1P)</math></li><li><math>\chi_{b2}(1P)</math></li><li><math>\eta_b(2S)</math></li><li><math>\Upsilon(2S)</math></li><li><math>\Upsilon(1D)</math></li><li><math>\chi_{b0}(2P)</math></li><li><math>\chi_{b1}(2P)</math></li><li><math>h_b(2P)</math></li><li><math>\chi_{b2}(2P)</math></li><li><math>\Upsilon(3S)</math></li><li><math>\chi_{b1}(3P)</math></li><li><math>\Upsilon(4S)</math></li><li><math>X(10610)^\pm</math></li><li><math>X(10610)^0</math></li><li><math>X(10650)^\pm</math></li><li><math>\Upsilon(10860)</math></li><li><math>\Upsilon(11020)</math></li></ul>	
				BOTTOM, CHARMED ( $B = C = \pm 1$ )		<ul style="list-style-type: none"><li><math>B_c^+</math></li><li><math>B_c(2S)^\pm</math></li></ul>			

## Baryon Summary Table

This short table gives the name, the quantum numbers (where known), and the status of baryons in the Review. Only the baryons with 3- or 4-star status are included in the Baryon Summary Table. Due to insufficient data or uncertain interpretation, the other entries in the table are not established baryons. The names with masses are of baryons that decay strongly. The spin-parity  $J^P$  (when known) is given with each particle. For the strongly decaying particles, the  $J^P$  values are considered to be part of the names.

$p$	$1/2^+$	****	$\Delta(1232)$	$3/2^+$	****	$\Sigma^+$	$1/2^+$	****	$\Xi^0$	$1/2^+$	****	$\Lambda_c^+$	$1/2^+$	****
$n$	$1/2^+$	****	$\Delta(1600)$	$3/2^+$	***	$\Sigma^0$	$1/2^+$	****	$\Xi^-$	$1/2^+$	****	$\Lambda_c(2595)^+$	$1/2^-$	***
$N(1440)$	$1/2^+$	****	$\Delta(1620)$	$1/2^-$	****	$\Sigma^-$	$1/2^+$	****	$\Xi(1530)$	$3/2^+$	****	$\Lambda_c(2625)^+$	$3/2^-$	***
$N(1520)$	$3/2^-$	****	$\Delta(1700)$	$3/2^-$	****	$\Sigma(1385)$	$3/2^+$	****	$\Xi(1620)$	*		$\Lambda_c(2765)^+$	*	
$N(1535)$	$1/2^-$	****	$\Delta(1750)$	$1/2^+$	*	$\Sigma(1480)$	*		$\Xi(1690)$	***		$\Lambda_c(2880)^+$	$5/2^+$	***
$N(1650)$	$1/2^-$	****	$\Delta(1900)$	$1/2^-$	**	$\Sigma(1560)$	**		$\Xi(1820)$	$3/2^-$	***	$\Lambda_c(2940)^+$	***	
$N(1675)$	$5/2^-$	****	$\Delta(1905)$	$5/2^+$	****	$\Sigma(1580)$	$3/2^-$	*	$\Xi(1950)$	***		$\Sigma_c(2455)$	$1/2^+$	****
$N(1680)$	$5/2^+$	****	$\Delta(1910)$	$1/2^+$	****	$\Sigma(1620)$	$1/2^-$	*	$\Xi(2030)$	$\geq \frac{5}{2}^?$	***	$\Sigma_c(2520)$	$3/2^+$	***
$N(1700)$	$3/2^-$	***	$\Delta(1920)$	$3/2^+$	***	$\Sigma(1660)$	$1/2^+$	***	$\Xi(2120)$	*		$\Sigma_c(2800)$	***	
$N(1710)$	$1/2^+$	****	$\Delta(1930)$	$5/2^-$	***	$\Sigma(1670)$	$3/2^-$	****	$\Xi(2250)$	**		$\Xi_c^+$	$1/2^+$	***
$N(1720)$	$3/2^+$	****	$\Delta(1940)$	$3/2^-$	**	$\Sigma(1690)$	**		$\Xi(2370)$	**		$\Xi_c^0$	$1/2^+$	***
$N(1860)$	$5/2^+$	**	$\Delta(1950)$	$7/2^+$	****	$\Sigma(1730)$	$3/2^+$	*	$\Xi(2500)$	*		$\Xi_c'^+$	$1/2^+$	***
$N(1875)$	$3/2^-$	***	$\Delta(2000)$	$5/2^+$	**	$\Sigma(1750)$	$1/2^-$	***				$\Xi_c'^0$	$1/2^+$	***
$N(1880)$	$1/2^+$	**	$\Delta(2150)$	$1/2^-$	*	$\Sigma(1770)$	$1/2^+$	*	$\Omega^-$	$3/2^+$	****	$\Xi_c$	$1/2^+$	***
$N(1895)$	$1/2^-$	**	$\Delta(2200)$	$7/2^-$	*	$\Sigma(1775)$	$5/2^-$	****	$\Omega(2250)^-$	***		$\Xi_c(2645)$	$3/2^+$	***
$N(1900)$	$3/2^+$	***	$\Delta(2300)$	$9/2^+$	**	$\Sigma(1840)$	$3/2^+$	*	$\Omega(2380)^-$	**		$\Xi_c(2790)$	$1/2^-$	***
$N(1990)$	$7/2^+$	**	$\Delta(2350)$	$5/2^-$	*	$\Sigma(1880)$	$1/2^+$	**	$\Omega(2470)^-$	**		$\Xi_c(2815)$	$3/2^-$	***
$N(2000)$	$5/2^+$	**	$\Delta(2390)$	$7/2^+$	*	$\Sigma(1900)$	$1/2^-$	*				$\Xi_c(2930)$	*	
$N(2040)$	$3/2^+$	*	$\Delta(2400)$	$9/2^-$	**	$\Sigma(1915)$	$5/2^+$	****				$\Xi_c(2970)$	***	
$N(2060)$	$5/2^-$	**	$\Delta(2420)$	$11/2^+$	****	$\Sigma(1940)$	$3/2^+$	*				$\Xi_c(3055)$	***	
$N(2100)$	$1/2^+$	*	$\Delta(2750)$	$13/2^-$	**	$\Sigma(1940)$	$3/2^-$	***				$\Xi_c(3080)$	***	
$N(2120)$	$3/2^-$	**	$\Delta(2950)$	$15/2^+$	**	$\Sigma(2000)$	$1/2^-$	*				$\Xi_c(3123)$	*	
$N(2190)$	$7/2^-$	****				$\Sigma(2030)$	$7/2^+$	****				$\Omega_c^0$	$1/2^+$	***
$N(2220)$	$9/2^+$	****	$\Lambda$	$1/2^+$	****	$\Sigma(2070)$	$5/2^+$	*				$\Omega_c(2770)^0$	$3/2^+$	***
$N(2250)$	$9/2^-$	****	$\Lambda(1405)$	$1/2^-$	****	$\Sigma(2080)$	$3/2^+$	**						
$N(2300)$	$1/2^+$	**	$\Lambda(1520)$	$3/2^-$	****	$\Sigma(2100)$	$7/2^-$	*				$\Xi_{cc}^+$	*	
$N(2570)$	$5/2^-$	**	$\Lambda(1600)$	$1/2^+$	***	$\Sigma(2250)$		***				$\Lambda_b^0$	$1/2^+$	***
$N(2600)$	$11/2^-$	***	$\Lambda(1670)$	$1/2^-$	****	$\Sigma(2455)$		**				$\Lambda_b(5912)^0$	$1/2^-$	***
$N(2700)$	$13/2^+$	**	$\Lambda(1690)$	$3/2^-$	****	$\Sigma(2620)$		**				$\Lambda_b(5920)^0$	$3/2^-$	***
			$\Lambda(1710)$	$1/2^+$	*	$\Sigma(3000)$		*				$\Sigma_b$	$1/2^+$	***
			$\Lambda(1800)$	$1/2^-$	***	$\Sigma(3170)$		*				$\Sigma_b^*$	$3/2^+$	***
			$\Lambda(1810)$	$1/2^+$	***							$\Xi_b^0, \Xi_b^-$	$1/2^+$	***
			$\Lambda(1820)$	$5/2^+$	****							$\Xi_b'(5935)^-$	$1/2^+$	***
			$\Lambda(1830)$	$5/2^-$	****							$\Xi_b(5945)^0$	$3/2^+$	***
			$\Lambda(1890)$	$3/2^+$	****							$\Xi_b^*(5955)^-$	$3/2^+$	***
			$\Lambda(2000)$	*								$\Omega_b^-$	$1/2^+$	***
			$\Lambda(2020)$	$7/2^+$	*									
			$\Lambda(2050)$	$3/2^-$	*							$P_c(4380)^+$	*	
			$\Lambda(2100)$	$7/2^-$	****							$P_c(4450)^+$	*	
			$\Lambda(2110)$	$5/2^+$	***									
			$\Lambda(2325)$	$3/2^-$	*									
			$\Lambda(2350)$	$9/2^+$	***									
			$\Lambda(2585)$	**										

\*\*\*\* Existence is certain, and properties are at least fairly well explored.

\*\*\* Existence ranges from very likely to certain, but further confirmation is desirable and/or quantum numbers, branching fractions, etc. are not well determined.

\*\* Evidence of existence is only fair.

\* Evidence of existence is poor.

# Baryon Summary Table

## $N$ BARYONS ( $S = 0$ , $I = 1/2$ )

$$p, N^+ = uud; \quad n, N^0 = udd$$

**p**

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

Mass  $m = 1.00727646688 \pm 0.00000000009$  u  
 Mass  $m = 938.272081 \pm 0.000006$  MeV <sup>[a]</sup>  
 $|m_p - m_{\bar{p}}|/m_p < 7 \times 10^{-10}$ , CL = 90% <sup>[b]</sup>  
 $|q_{\bar{p}}|/(q_p) = 0.9999999991 \pm 0.0000000009$   
 $|q_p + q_{\bar{p}}|/e < 7 \times 10^{-10}$ , CL = 90% <sup>[b]</sup>  
 $|q_p + q_e|/e < 1 \times 10^{-21}$  [c]  
 Magnetic moment  $\mu = 2.792847351 \pm 0.000000009$   $\mu_N$   
 $(\mu_p + \mu_{\bar{p}})/\mu_p = (0 \pm 5) \times 10^{-6}$   
 Electric dipole moment  $d < 0.54 \times 10^{-23}$  ecm  
 Electric polarizability  $\alpha = (11.2 \pm 0.4) \times 10^{-4}$  fm<sup>3</sup>  
 Magnetic polarizability  $\beta = (2.5 \pm 0.4) \times 10^{-4}$  fm<sup>3</sup> ( $S = 1.2$ )  
 Charge radius,  $\mu p$  Lamb shift =  $0.84087 \pm 0.00039$  fm <sup>[d]</sup>  
 Charge radius,  $e p$  CODATA value =  $0.8751 \pm 0.0061$  fm <sup>[d]</sup>  
 Magnetic radius =  $0.78 \pm 0.04$  fm <sup>[e]</sup>  
 Mean life  $\tau > 2.1 \times 10^{29}$  years, CL = 90% <sup>[f]</sup> ( $p \rightarrow$  invisible mode)  
 Mean life  $\tau > 10^{31}$  to  $10^{33}$  years <sup>[f]</sup> (mode dependent)

See the "Note on Nucleon Decay" in our 1994 edition (Phys. Rev. **D50**, 1173) for a short review.

The "partial mean life" limits tabulated here are the limits on  $\tau/B_i$ , where  $\tau$  is the total mean life and  $B_i$  is the branching fraction for the mode in question. For  $N$  decays,  $p$  and  $n$  indicate proton and neutron partial lifetimes.

$p$ DECAY MODES	Partial mean life ( $10^{30}$ years)	Confidence level	$p$ (MeV/c)
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### Antilepton + meson

$N \rightarrow e^+ \pi$	$> 2000$ ( $n$ ), $> 8200$ ( $p$ )	90%	459
$N \rightarrow \mu^+ \pi$	$> 1000$ ( $n$ ), $> 6600$ ( $p$ )	90%	453
$N \rightarrow \nu \pi$	$> 1100$ ( $n$ ), $> 390$ ( $p$ )	90%	459
$p \rightarrow e^+ \eta$	$> 4200$	90%	309
$p \rightarrow \mu^+ \eta$	$> 1300$	90%	297
$n \rightarrow \nu \eta$	$> 158$	90%	310
$N \rightarrow e^+ \rho$	$> 217$ ( $n$ ), $> 710$ ( $p$ )	90%	149
$N \rightarrow \mu^+ \rho$	$> 228$ ( $n$ ), $> 160$ ( $p$ )	90%	113
$N \rightarrow \nu \rho$	$> 19$ ( $n$ ), $> 162$ ( $p$ )	90%	149
$p \rightarrow e^+ \omega$	$> 320$	90%	143
$p \rightarrow \mu^+ \omega$	$> 780$	90%	105
$n \rightarrow \nu \omega$	$> 108$	90%	144
$N \rightarrow e^+ K$	$> 17$ ( $n$ ), $> 1000$ ( $p$ )	90%	339
$N \rightarrow \mu^+ K$	$> 26$ ( $n$ ), $> 1600$ ( $p$ )	90%	329
$N \rightarrow \nu K$	$> 86$ ( $n$ ), $> 5900$ ( $p$ )	90%	339
$n \rightarrow \nu K_S^0$	$> 260$	90%	338
$p \rightarrow e^+ K^*(892)^0$	$> 84$	90%	45
$N \rightarrow \nu K^*(892)$	$> 78$ ( $n$ ), $> 51$ ( $p$ )	90%	45

### Antilepton + mesons

$p \rightarrow e^+ \pi^+ \pi^-$	$> 82$	90%	448
$p \rightarrow e^+ \pi^0 \pi^0$	$> 147$	90%	449
$n \rightarrow e^+ \pi^- \pi^0$	$> 52$	90%	449
$p \rightarrow \mu^+ \pi^+ \pi^-$	$> 133$	90%	425
$p \rightarrow \mu^+ \pi^0 \pi^0$	$> 101$	90%	427
$n \rightarrow \mu^+ \pi^- \pi^0$	$> 74$	90%	427
$n \rightarrow e^+ K^0 \pi^-$	$> 18$	90%	319

### Lepton + meson

$n \rightarrow e^- \pi^+$	$> 65$	90%	459
$n \rightarrow \mu^- \pi^+$	$> 49$	90%	453
$n \rightarrow e^- \rho^+$	$> 62$	90%	150
$n \rightarrow \mu^- \rho^+$	$> 7$	90%	115
$n \rightarrow e^- K^+$	$> 32$	90%	340
$n \rightarrow \mu^- K^+$	$> 57$	90%	330

### Lepton + mesons

$p \rightarrow e^- \pi^+ \pi^+$	$> 30$	90%	448
$n \rightarrow e^- \pi^+ \pi^0$	$> 29$	90%	449
$p \rightarrow \mu^- \pi^+ \pi^+$	$> 17$	90%	425
$n \rightarrow \mu^- \pi^+ \pi^0$	$> 34$	90%	427
$p \rightarrow e^- \pi^+ K^+$	$> 75$	90%	320
$p \rightarrow \mu^- \pi^+ K^+$	$> 245$	90%	279

### Antilepton + photon(s)

$p \rightarrow e^+ \gamma$	$> 670$	90%	469
$p \rightarrow \mu^+ \gamma$	$> 478$	90%	463
$n \rightarrow \nu \gamma$	$> 550$	90%	470
$p \rightarrow e^+ \gamma \gamma$	$> 100$	90%	469
$n \rightarrow \nu \gamma \gamma$	$> 219$	90%	470

### Antilepton + single massless

$p \rightarrow e^+ X$	$> 790$	90%	—
$p \rightarrow \mu^+ X$	$> 410$	90%	—

### Three (or more) leptons

$p \rightarrow e^+ e^+ e^-$	$> 793$	90%	469
$p \rightarrow e^+ \mu^+ \mu^-$	$> 359$	90%	457
$p \rightarrow e^+ \nu \nu$	$> 170$	90%	469
$n \rightarrow e^+ e^- \nu$	$> 257$	90%	470
$n \rightarrow \mu^+ e^- \nu$	$> 83$	90%	464
$n \rightarrow \mu^+ \mu^- \nu$	$> 79$	90%	458
$p \rightarrow \mu^+ e^+ e^-$	$> 529$	90%	463
$p \rightarrow \mu^+ \mu^+ \mu^-$	$> 675$	90%	439
$p \rightarrow \mu^+ \nu \nu$	$> 220$	90%	463
$p \rightarrow e^- \mu^+ \mu^+$	$> 6$	90%	457
$n \rightarrow 3 \nu$	$> 5 \times 10^{-4}$	90%	470

### Inclusive modes

$N \rightarrow e^+$ anything	$> 0.6$ ( $n, p$ )	90%	—
$N \rightarrow \mu^+$ anything	$> 12$ ( $n, p$ )	90%	—
$N \rightarrow e^+ \pi^0$ anything	$> 0.6$ ( $n, p$ )	90%	—

### $\Delta B = 2$ dinucleon modes

The following are lifetime limits per iron nucleus.

$pp \rightarrow \pi^+ \pi^+$	$> 72.2$	90%	—
$p n \rightarrow \pi^+ \pi^0$	$> 170$	90%	—
$n n \rightarrow \pi^+ \pi^-$	$> 0.7$	90%	—
$n n \rightarrow \pi^0 \pi^0$	$> 404$	90%	—
$pp \rightarrow K^+ K^+$	$> 170$	90%	—
$pp \rightarrow e^+ e^+$	$> 5.8$	90%	—
$pp \rightarrow e^+ \mu^+$	$> 3.6$	90%	—
$pp \rightarrow \mu^+ \mu^+$	$> 1.7$	90%	—
$p n \rightarrow e^+ \bar{\nu}$	$> 260$	90%	—
$p n \rightarrow \mu^+ \bar{\nu}$	$> 200$	90%	—
$p n \rightarrow \tau^+ \bar{\nu}_\tau$	$> 29$	90%	—
$n n \rightarrow \nu_e \bar{\nu}_e$	$> 1.4$	90%	—
$n n \rightarrow \nu_\mu \bar{\nu}_\mu$	$> 1.4$	90%	—
$p n \rightarrow$ invisible	$> 2.1 \times 10^{-5}$	90%	—
$pp \rightarrow$ invisible	$> 5 \times 10^{-5}$	90%	—

### $\bar{p}$ DECAY MODES

$\bar{p}$ DECAY MODES	Partial mean life (years)	Confidence level	$p$ (MeV/c)
$\bar{p} \rightarrow e^- \gamma$	$> 7 \times 10^5$	90%	469
$\bar{p} \rightarrow \mu^- \gamma$	$> 5 \times 10^4$	90%	463
$\bar{p} \rightarrow e^- \pi^0$	$> 4 \times 10^5$	90%	459
$\bar{p} \rightarrow \mu^- \pi^0$	$> 5 \times 10^4$	90%	453
$\bar{p} \rightarrow e^- \eta$	$> 2 \times 10^4$	90%	309
$\bar{p} \rightarrow \mu^- \eta$	$> 8 \times 10^3$	90%	297
$\bar{p} \rightarrow e^- K_S^0$	$> 900$	90%	337
$\bar{p} \rightarrow \mu^- K_S^0$	$> 4 \times 10^3$	90%	326
$\bar{p} \rightarrow e^- K_L^0$	$> 9 \times 10^3$	90%	337
$\bar{p} \rightarrow \mu^- K_L^0$	$> 7 \times 10^3$	90%	326
$\bar{p} \rightarrow e^- \gamma \gamma$	$> 2 \times 10^4$	90%	469
$\bar{p} \rightarrow \mu^- \gamma \gamma$	$> 2 \times 10^4$	90%	463
$\bar{p} \rightarrow e^- \omega$	$> 200$	90%	143

**n**

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

Mass  $m = 1.0086649159 \pm 0.0000000005$  u  
 Mass  $m = 939.565413 \pm 0.000006$  MeV <sup>[a]</sup>  
 $(m_n - m_{\bar{n}})/m_n = (9 \pm 6) \times 10^{-5}$   
 $m_n - m_p = 1.2933321 \pm 0.0000005$  MeV  
 $= 0.00138844919(45)$  u  
 Mean life  $\tau = 880.2 \pm 1.0$  s ( $S = 1.9$ )  
 $c\tau = 2.6387 \times 10^8$  km  
 Magnetic moment  $\mu = -1.9130427 \pm 0.0000005$   $\mu_N$   
 Electric dipole moment  $d < 0.30 \times 10^{-25}$  ecm, CL = 90%  
 Mean-square charge radius  $\langle r_n^2 \rangle = -0.1161 \pm 0.0022$  fm<sup>2</sup> ( $S = 1.3$ )



## Baryon Summary Table

Magnetic radius  $\sqrt{\langle r_M^2 \rangle} = 0.864^{+0.009}_{-0.008}$  fm  
 Electric polarizability  $\alpha = (11.8 \pm 1.1) \times 10^{-4}$  fm<sup>3</sup>  
 Magnetic polarizability  $\beta = (3.7 \pm 1.2) \times 10^{-4}$  fm<sup>3</sup>  
 Charge  $q = (-0.2 \pm 0.8) \times 10^{-21}$  e  
 Mean  $n\bar{\pi}$ -oscillation time  $> 2.7 \times 10^8$  s, CL = 90% (free  $n$ )  
 Mean  $n\bar{\pi}$ -oscillation time  $> 1.3 \times 10^8$  s, CL = 90% [g] (bound  $n$ )  
 Mean  $n\eta'$ -oscillation time  $> 414$  s, CL = 90% [h]

 **$p e^- \nu_e$  decay parameters [l]**

$\lambda \equiv g_A / g_V = -1.2723 \pm 0.0023$  (S = 2.2)  
 $A = -0.1184 \pm 0.0010$  (S = 2.4)  
 $B = 0.9807 \pm 0.0030$   
 $C = -0.2377 \pm 0.0026$   
 $a = -0.103 \pm 0.004$   
 $\phi_{AV} = (180.017 \pm 0.026)^\circ$  [l]  
 $D = (-1.2 \pm 2.0) \times 10^{-4}$  [k]  
 $R = 0.004 \pm 0.013$  [k]

<b><math>n</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$p e^- \bar{\nu}_e$	100	%	1
$p e^- \bar{\nu}_e \gamma$	[l] $(3.09 \pm 0.32) \times 10^{-3}$		1
<b>Charge conservation (Q) violating mode</b>			
$p \nu_e \bar{\nu}_e$	Q < 8	$\times 10^{-27}$ 68%	1

 **$N(1440) 1/2^+$** 

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

Re(pole position)  
 $-2\text{Im}(\text{pole position}) = 100$  to 170 ( $\approx 135$ ) MeV  
 Breit-Wigner mass = 1410 to 1450 ( $\approx 1430$ ) MeV  
 Breit-Wigner full width = 250 to 450 ( $\approx 350$ ) MeV

<b><math>N(1440)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	55–75 %	391
$N\eta$	<1 %	†
$N\pi\pi$	25–50 %	338
$\Delta(1232)\pi$	20–30 %	135
$\Delta(1232)\pi, P\text{-wave}$	13–27 %	135
$N\sigma$	11–23 %	–
$p\gamma, \text{helicity}=1/2$	0.035–0.048 %	407
$n\gamma, \text{helicity}=1/2$	0.02–0.04 %	406

 **$N(1520) 3/2^-$** 

$$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$$

Re(pole position) = 1505 to 1515 ( $\approx 1510$ ) MeV  
 $-2\text{Im}(\text{pole position}) = 105$  to 120 ( $\approx 110$ ) MeV  
 Breit-Wigner mass = 1510 to 1520 ( $\approx 1515$ ) MeV  
 Breit-Wigner full width = 100 to 125 ( $\approx 115$ ) MeV

<b><math>N(1520)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	55–65 %	453
$N\eta$	<1 %	142
$N\pi\pi$	25–35 %	410
$\Delta(1232)\pi$	22–34 %	225
$\Delta(1232)\pi, S\text{-wave}$	15–23 %	225
$\Delta(1232)\pi, D\text{-wave}$	7–11 %	225
$N\sigma$	<2 %	–
$p\gamma$	0.31–0.52 %	467
$p\gamma, \text{helicity}=1/2$	0.01–0.02 %	467
$p\gamma, \text{helicity}=3/2$	0.30–0.50 %	467
$n\gamma$	0.30–0.53 %	466
$n\gamma, \text{helicity}=1/2$	0.04–0.10 %	466
$n\gamma, \text{helicity}=3/2$	0.25–0.45 %	466

 **$N(1535) 1/2^-$** 

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^-)$$

Re(pole position) = 1490 to 1530 ( $\approx 1510$ ) MeV  
 $-2\text{Im}(\text{pole position}) = 90$  to 250 ( $\approx 170$ ) MeV  
 Breit-Wigner mass = 1525 to 1545 ( $\approx 1535$ ) MeV  
 Breit-Wigner full width = 125 to 175 ( $\approx 150$ ) MeV

<b><math>N(1535)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	35–55 %	468
$N\eta$	32–52 %	186
$N\pi\pi$	3–14 %	426
$\Delta(1232)\pi, D\text{-wave}$	1–4 %	244
$N\sigma$	2–10 %	–
$N(1440)\pi$	5–12 %	†
$p\gamma, \text{helicity}=1/2$	0.15–0.30 %	481
$n\gamma, \text{helicity}=1/2$	0.01–0.25 %	480

 **$N(1650) 1/2^-$** 

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^-)$$

Re(pole position) = 1640 to 1670 ( $\approx 1655$ ) MeV  
 $-2\text{Im}(\text{pole position}) = 100$  to 170 ( $\approx 135$ ) MeV  
 Breit-Wigner mass = 1645 to 1670 ( $\approx 1655$ ) MeV  
 Breit-Wigner full width = 110 to 170 ( $\approx 140$ ) MeV

<b><math>N(1650)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	50–70 %	551
$N\eta$	14–22 %	354
$\Lambda K$	5–15 %	179
$N\pi\pi$	8–36 %	517
$\Delta(1232)\pi, D\text{-wave}$	6–18 %	349
$N\sigma$	2–18 %	–
$N(1440)\pi$	6–26 %	168
$p\gamma, \text{helicity}=1/2$	0.04–0.20 %	562
$n\gamma, \text{helicity}=1/2$	0.003–0.17 %	561

 **$N(1675) 5/2^-$** 

$$I(J^P) = \frac{1}{2}(\frac{5}{2}^-)$$

Re(pole position) = 1655 to 1665 ( $\approx 1660$ ) MeV  
 $-2\text{Im}(\text{pole position}) = 125$  to 150 ( $\approx 135$ ) MeV  
 Breit-Wigner mass = 1670 to 1680 ( $\approx 1675$ ) MeV  
 Breit-Wigner full width = 130 to 165 ( $\approx 150$ ) MeV

<b><math>N(1675)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	35–45 %	564
$N\eta$	<1 %	376
$N\pi\pi$	25–45 %	532
$\Delta(1232)\pi, D\text{-wave}$	23–37 %	366
$N\sigma$	3–7 %	–
$p\gamma$	0–0.02 %	575
$p\gamma, \text{helicity}=1/2$	0–0.01 %	575
$p\gamma, \text{helicity}=3/2$	0–0.01 %	575
$n\gamma$	0–0.15 %	574
$n\gamma, \text{helicity}=1/2$	0–0.05 %	574
$n\gamma, \text{helicity}=3/2$	0–0.10 %	574

 **$N(1680) 5/2^+$** 

$$I(J^P) = \frac{1}{2}(\frac{5}{2}^+)$$

Re(pole position) = 1665 to 1680 ( $\approx 1675$ ) MeV  
 $-2\text{Im}(\text{pole position}) = 110$  to 135 ( $\approx 120$ ) MeV  
 Breit-Wigner mass = 1680 to 1690 ( $\approx 1685$ ) MeV  
 Breit-Wigner full width = 120 to 140 ( $\approx 130$ ) MeV

<b><math>N(1680)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	65–70 %	571
$N\eta$	<1 %	386
$N\pi\pi$	20–40 %	539
$\Delta(1232)\pi$	11–23 %	374
$\Delta(1232)\pi, P\text{-wave}$	4–10 %	374
$\Delta(1232)\pi, F\text{-wave}$	7–13 %	374
$N\sigma$	9–19 %	–
$p\gamma$	0.21–0.32 %	581
$p\gamma, \text{helicity}=1/2$	0.001–0.011 %	581
$p\gamma, \text{helicity}=3/2$	0.20–0.32 %	581
$n\gamma$	0.021–0.046 %	581
$n\gamma, \text{helicity}=1/2$	0.004–0.029 %	581
$n\gamma, \text{helicity}=3/2$	0.01–0.024 %	581

# Baryon Summary Table

**$N(1700) \ 3/2^-$**

$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$

Re(pole position) = 1650 to 1750 ( $\approx$  1700) MeV  
−2Im(pole position) = 100 to 300 MeV  
Breit-Wigner mass = 1650 to 1750 ( $\approx$  1700) MeV  
Breit-Wigner full width = 100 to 250 ( $\approx$  150) MeV

<b><math>N(1700)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	7–17 %	581
$N\eta$	seen	402
$N\pi\pi$	60–90 %	550
$\Delta(1232)\pi$	55–85 %	386
$\Delta(1232)\pi$ , $S$ -wave	50–80 %	386
$\Delta(1232)\pi$ , $D$ -wave	4–14 %	386
$N(1440)\pi$	3–11 %	215
$N(1520)\pi$	<4 %	120
$N\rho$ , $S=3/2$ , $S$ -wave	seen	†
$N\sigma$	2–14 %	–
$p\gamma$	0.01–0.05 %	591
$p\gamma$ , helicity=1/2	0.0–0.024 %	591
$p\gamma$ , helicity=3/2	0.002–0.026 %	591
$n\gamma$	0.01–0.13 %	590
$n\gamma$ , helicity=1/2	0.0–0.09 %	590
$n\gamma$ , helicity=3/2	0.01–0.05 %	590

**$N(1710) \ 1/2^+$**

$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$

Re(pole position) = 1670 to 1770 ( $\approx$  1720) MeV  
−2Im(pole position) = 80 to 380 ( $\approx$  230) MeV  
Breit-Wigner mass = 1680 to 1740 ( $\approx$  1710) MeV  
Breit-Wigner full width = 50 to 250 ( $\approx$  100) MeV

<b><math>N(1710)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	5–20 %	588
$N\eta$	10–50 %	412
$N\omega$	1–5 %	†
$\Lambda K$	5–25 %	269
$\Sigma K$	seen	138
$N\pi\pi$	seen	557
$\Delta(1232)\pi$ , $P$ -wave	seen	394
$N(1535)\pi$	9–21 %	106
$N\rho$ , $S=1/2$ , $P$ -wave	seen	†
$p\gamma$ , helicity=1/2	0.002–0.08 %	598
$n\gamma$ , helicity=1/2	0.0–0.02%	597

**$N(1720) \ 3/2^+$**

$I(J^P) = \frac{1}{2}(\frac{3}{2}^+)$

Re(pole position) = 1660 to 1690 ( $\approx$  1675) MeV  
−2Im(pole position) = 150 to 400 ( $\approx$  250) MeV  
Breit-Wigner mass = 1700 to 1750 ( $\approx$  1720) MeV  
Breit-Wigner full width = 150 to 400 ( $\approx$  250) MeV

<b><math>N(1720)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	8–14 %	594
$N\eta$	1–5 %	422
$\Lambda K$	4–5 %	283
$N\pi\pi$	50–90 %	564
$\Delta(1232)\pi$ , $P$ -wave	47–77 %	402
$\Delta(1232)\pi$ , $F$ -wave	<12 %	402
$N\rho$	70–85 %	74
$N\rho$ , $S=1/2$ , $P$ -wave	seen	74
$N\sigma$	2–14 %	–
$N(1440)\pi$	<2 %	235
$N(1520)\pi$ , $S$ -wave	1–5 %	145
$p\gamma$	0.05–0.25 %	604
$p\gamma$ , helicity=1/2	0.05–0.15 %	604
$p\gamma$ , helicity=3/2	0.002–0.16 %	604
$n\gamma$	0.0–0.016 %	603
$n\gamma$ , helicity=1/2	0.0–0.01 %	603
$n\gamma$ , helicity=3/2	0.0–0.015 %	603

**$N(1875) \ 3/2^-$**

$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$

Re(pole position) = 1800 to 1950 MeV  
−2Im(pole position) = 150 to 250 MeV  
Breit-Wigner mass = 1820 to 1920 ( $\approx$  1875) MeV  
Breit-Wigner full width = 250  $\pm$  70 MeV

<b><math>N(1875)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	2–14 %	695
$N\eta$	<1 %	559
$N\omega$	15–25 %	371
$\Lambda K$	seen	454
$\Sigma K$	seen	384
$N\pi\pi$		670
$\Delta(1232)\pi$	10–35 %	520
$\Delta(1232)\pi$ , $S$ -wave	7–21 %	520
$\Delta(1232)\pi$ , $D$ -wave	2–12 %	520
$N\rho$ , $S=3/2$ , $S$ -wave	seen	379
$N\sigma$	30–60 %	–
$N(1440)\pi$	2–8 %	373
$N(1520)\pi$	<2 %	301
$p\gamma$	0.001–0.025 %	703
$p\gamma$ , helicity=1/2	0.001–0.021 %	703
$p\gamma$ , helicity=3/2	<0.003 %	703
$n\gamma$	<0.040 %	702
$n\gamma$ , helicity=1/2	<0.007 %	702
$n\gamma$ , helicity=3/2	<0.033 %	702

**$N(1900) \ 3/2^+$**

$I(J^P) = \frac{1}{2}(\frac{3}{2}^+)$

Re(pole position) = 1900 to 1940 ( $\approx$  1920) MeV  
−2Im(pole position) = 130 to 300 MeV  
Breit-Wigner mass = 1900  $\pm$  30 MeV  
Breit-Wigner full width = 200  $\pm$  50 MeV

<b><math>N(1900)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	<10 %	710
$N\eta$	2–14 %	579
$N\omega$	7–13 %	401
$\Lambda K$	2–20 %	477
$\Sigma K$	3–7 %	410
$N\pi\pi$	40–80 %	686
$\Delta(1232)\pi$	30–70 %	539
$\Delta(1232)\pi$ , $P$ -wave	9–25 %	539
$\Delta(1232)\pi$ , $F$ -wave	21–45 %	539
$N\sigma$	1–7 %	–
$N(1520)\pi$	7–23 %	324
$N(1535)\pi$	4–10 %	306
$p\gamma$	0.001–0.025 %	718
$p\gamma$ , helicity=1/2	0.001–0.021 %	718
$p\gamma$ , helicity=3/2	<0.003 %	718
$n\gamma$	<0.040 %	718
$n\gamma$ , helicity=1/2	<0.007 %	718
$n\gamma$ , helicity=3/2	<0.033 %	718

**$N(2190) \ 7/2^-$**

$I(J^P) = \frac{1}{2}(\frac{7}{2}^-)$

Re(pole position) = 2050 to 2100 ( $\approx$  2075) MeV  
−2Im(pole position) = 400 to 520 ( $\approx$  450) MeV  
Breit-Wigner mass = 2100 to 2200 ( $\approx$  2190) MeV  
Breit-Wigner full width = 300 to 700 ( $\approx$  500) MeV

<b><math>N(2190)</math> DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
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## Baryon Summary Table

<b>N(2190) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	10–20 %	888
$N\eta$	seen	791
$\Lambda K$	0.2–0.8;%	712
$N\pi\pi$	22–80;%	870
$\Delta(1232)\pi$ , $D$ -wave	19–31 %	740
$N\rho$ , $S=3/2$ , $D$ -wave	seen	680
$N\sigma$	3–9 %	–
$p\gamma$	0.014–0.077 %	894
$p\gamma$ , helicity=1/2	0.013–0.062;%	894
$p\gamma$ , helicity=3/2	0.001–0.014;%	894
$n\gamma$	<0.04 %	893
$n\gamma$ , helicity=1/2	<0.01;%	893
$n\gamma$ , helicity=3/2	<0.03 %	893

$$\boxed{N(2220) \ 9/2^+} \quad I(J^P) = \frac{1}{2}(\frac{9}{2}^+)$$

Re(pole position) = 2130 to 2200 ( $\approx$  2170) MeV  
 $-2\text{Im}(\text{pole position}) = 400$  to 560 ( $\approx$  480) MeV  
 Breit-Wigner mass = 2200 to 2300 ( $\approx$  2250) MeV  
 Breit-Wigner full width = 350 to 500 ( $\approx$  400) MeV

<b>N(2220) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	15–25 %	924

$$\boxed{N(2250) \ 9/2^-} \quad I(J^P) = \frac{1}{2}(\frac{9}{2}^-)$$

Re(pole position) = 2150 to 2250 ( $\approx$  2200) MeV  
 $-2\text{Im}(\text{pole position}) = 350$  to 550 ( $\approx$  450) MeV  
 Breit-Wigner mass = 2250 to 2320 ( $\approx$  2280) MeV  
 Breit-Wigner full width = 300 to 600 ( $\approx$  500) MeV

<b>N(2250) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	5–15 %	941

$$\boxed{N(2600) \ 11/2^-} \quad I(J^P) = \frac{1}{2}(\frac{11}{2}^-)$$

Breit-Wigner mass = 2550 to 2750 ( $\approx$  2600) MeV  
 Breit-Wigner full width = 500 to 800 ( $\approx$  650) MeV

<b>N(2600) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	5–10 %	1126

## Δ BARYONS ( $S=0$ , $I=3/2$ )

$$\Delta^{++} = uuu, \quad \Delta^+ = uud, \quad \Delta^0 = udd, \quad \Delta^- = ddd$$

$$\boxed{\Delta(1232) \ 3/2^+} \quad I(J^P) = \frac{3}{2}(\frac{3}{2}^+)$$

Re(pole position) = 1209 to 1211 ( $\approx$  1210) MeV  
 $-2\text{Im}(\text{pole position}) = 98$  to 102 ( $\approx$  100) MeV  
 Breit-Wigner mass (mixed charges) = 1230 to 1234 ( $\approx$  1232) MeV  
 Breit-Wigner full width (mixed charges) = 114 to 120 ( $\approx$  117) MeV

<b>Δ(1232) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	99.4 %	229
$N\gamma$	0.55–0.65 %	259
$N\gamma$ , helicity=1/2	0.11–0.13 %	259
$N\gamma$ , helicity=3/2	0.44–0.52 %	259

$$\boxed{\Delta(1600) \ 3/2^+} \quad I(J^P) = \frac{3}{2}(\frac{3}{2}^+)$$

Re(pole position) = 1460 to 1560 ( $\approx$  1510) MeV  
 $-2\text{Im}(\text{pole position}) = 200$  to 350 ( $\approx$  275) MeV  
 Breit-Wigner mass = 1500 to 1700 ( $\approx$  1600) MeV  
 Breit-Wigner full width = 220 to 420 ( $\approx$  320) MeV

<b>Δ(1600) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	10–25 %	513
$N\pi\pi$	75–90 %	477
$\Delta(1232)\pi$	73–83 %	303
$\Delta(1232)\pi$ , $P$ -wave	72–82 %	303
$\Delta(1232)\pi$ , $F$ -wave	<2 %	303
$N(1440)\pi$ , $P$ -wave	seen	98
$N\gamma$	0.001–0.035 %	525
$N\gamma$ , helicity=1/2	0.0–0.02 %	525
$N\gamma$ , helicity=3/2	0.001–0.015 %	525

$$\boxed{\Delta(1620) \ 1/2^-} \quad I(J^P) = \frac{3}{2}(\frac{1}{2}^-)$$

Re(pole position) = 1590 to 1610 ( $\approx$  1600) MeV  
 $-2\text{Im}(\text{pole position}) = 120$  to 140 ( $\approx$  130) MeV  
 Breit-Wigner mass = 1600 to 1660 ( $\approx$  1630) MeV  
 Breit-Wigner full width = 130 to 150 ( $\approx$  140) MeV

<b>Δ(1620) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	20–30 %	534
$N\pi\pi$	55–80 %	499
$\Delta(1232)\pi$ , $D$ -wave	52–72 %	328
$N\rho$ , $S=1/2$ , $S$ -wave	seen	†
$N\rho$ , $S=3/2$ , $D$ -wave	seen	†
$N(1440)\pi$	3–9 %	138
$N\gamma$ , helicity=1/2	0.03–0.10 %	545

$$\boxed{\Delta(1700) \ 3/2^-} \quad I(J^P) = \frac{3}{2}(\frac{3}{2}^-)$$

Re(pole position) = 1620 to 1680 ( $\approx$  1650) MeV  
 $-2\text{Im}(\text{pole position}) = 160$  to 300 ( $\approx$  230) MeV  
 Breit-Wigner mass = 1670 to 1750 ( $\approx$  1700) MeV  
 Breit-Wigner full width = 200 to 400 ( $\approx$  300) MeV

<b>Δ(1700) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	10–20 %	581
$N\pi\pi$	10–55 %	550
$\Delta(1232)\pi$	10–50 %	386
$\Delta(1232)\pi$ , $S$ -wave	5–35 %	386
$\Delta(1232)\pi$ , $D$ -wave	4–16 %	386
$N\rho$ , $S=3/2$ , $S$ -wave	seen	†
$N(1520)\pi$ , $P$ -wave	1–5 %	120
$N(1535)\pi$	0.5–1.5 %	90
$\Delta(1232)\eta$	3–7 %	†
$N\gamma$	0.22–0.60 %	591
$N\gamma$ , helicity=1/2	0.12–0.30 %	591
$N\gamma$ , helicity=3/2	0.10–0.30 %	591

$$\boxed{\Delta(1905) \ 5/2^+} \quad I(J^P) = \frac{3}{2}(\frac{5}{2}^+)$$

Re(pole position) = 1805 to 1835 ( $\approx$  1820) MeV  
 $-2\text{Im}(\text{pole position}) = 265$  to 300 ( $\approx$  280) MeV  
 Breit-Wigner mass = 1855 to 1910 ( $\approx$  1880) MeV  
 Breit-Wigner full width = 270 to 400 ( $\approx$  330) MeV

<b>Δ(1905) DECAY MODES</b>	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\pi$	9–15 %	698
$N\pi\pi$		673
$\Delta(1232)\pi$ , $P$ -wave	23–43 %	524
$\Delta(1232)\pi$ , $F$ -wave	seen	524
$N\rho$ , $S=3/2$ , $P$ -wave	seen	385
$N(1535)\pi$	< 1 %	288
$N(1680)\pi$ , $P$ -wave	5–15 %	133
$\Delta(1232)\eta$	2–6 %	282
$N\gamma$	0.012–0.036 %	706
$N\gamma$ , helicity=1/2	0.002–0.006 %	706
$N\gamma$ , helicity=3/2	0.01–0.03 %	706

$$\boxed{\Delta(1910) \ 1/2^+} \quad I(J^P) = \frac{3}{2}(\frac{1}{2}^+)$$

Re(pole position) = 1830 to 1880 ( $\approx$  1855) MeV  
 $-2\text{Im}(\text{pole position}) = 200$  to 500 ( $\approx$  350) MeV  
 Breit-Wigner mass = 1860 to 1910 ( $\approx$  1890) MeV  
 Breit-Wigner full width = 220 to 340 ( $\approx$  280) MeV

# Baryon Summary Table

$\Delta(1910)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\pi$	15–30 %	704
$\Sigma K$	4–14 %	400
$N\pi\pi$		680
$\Delta(1232)\pi$	34–66 %	531
$N(1440)\pi$	3–9 %	386
$\Delta(1232)\eta$	5–13 %	296
$N\gamma$ , helicity=1/2	0.0–0.02 %	712

$\Delta(1920) 3/2^+$

$I(J^P) = \frac{3}{2}(\frac{3}{2}^+)$

Re(pole position) = 1850 to 1950 ( $\approx$  1900) MeV  
–2Im(pole position) = 200 to 400 ( $\approx$  300) MeV  
Breit-Wigner mass = 1900 to 1970 ( $\approx$  1920) MeV  
Breit-Wigner full width = 180 to 300 ( $\approx$  260) MeV

$\Delta(1920)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\pi$	5–20 %	723
$\Sigma K$	2–6 %	431
$N\pi\pi$		699
$\Delta(1232)\pi$	50–90 %	553
$\Delta(1232)\pi$ , $P$ -wave	8–28 %	553
$\Delta(1232)\pi$ , $F$ -wave	44–72 %	553
$N(1440)\pi$ , $P$ -wave	<4 %	411
$N(1520)\pi$ , $S$ -wave	<5 %	341
$N(1535)\pi$	<2 %	324
$N\pi_0(980)$	seen	41
$\Delta(1232)\eta$	5–17 %	336

$\Delta(1930) 5/2^-$

$I(J^P) = \frac{3}{2}(\frac{5}{2}^-)$

Re(pole position) = 1840 to 1960 ( $\approx$  1900) MeV  
–2Im(pole position) = 175 to 360 ( $\approx$  270) MeV  
Breit-Wigner mass = 1900 to 2000 ( $\approx$  1950) MeV  
Breit-Wigner full width = 220 to 500 ( $\approx$  360) MeV

$\Delta(1930)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\pi$	5–15 %	742
$N\gamma$	0.0–0.01 %	749
$N\gamma$ , helicity=1/2	0.0–0.005 %	749
$N\gamma$ , helicity=3/2	0.0–0.004 %	749

$\Delta(1950) 7/2^+$

$I(J^P) = \frac{3}{2}(\frac{7}{2}^+)$

Re(pole position) = 1870 to 1890 ( $\approx$  1880) MeV  
–2Im(pole position) = 220 to 260 ( $\approx$  240) MeV  
Breit-Wigner mass = 1915 to 1950 ( $\approx$  1930) MeV  
Breit-Wigner full width = 235 to 335 ( $\approx$  285) MeV

$\Delta(1950)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\pi$	35–45 %	729
$\Sigma K$	0.3–0.5 %	441
$N\pi\pi$		706
$\Delta(1232)\pi$ , $F$ -wave	1–9 %	560
$N(1680)\pi$ , $P$ -wave	3–9 %	191
$\Delta(1232)\eta$	< 1 %	349

$\Delta(2420) 11/2^+$

$I(J^P) = \frac{3}{2}(\frac{11}{2}^+)$

Re(pole position) = 2260 to 2400 ( $\approx$  2330) MeV  
–2Im(pole position) = 350 to 750 ( $\approx$  550) MeV  
Breit-Wigner mass = 2300 to 2500 ( $\approx$  2420) MeV  
Breit-Wigner full width = 300 to 500 ( $\approx$  400) MeV

$\Delta(2420)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\pi$	5–15 %	1023

$\Lambda$  BARYONS  
( $S = -1$ ,  $I = 0$ )  
 $\Lambda^0 = uds$

$\Lambda$

$I(J^P) = 0(\frac{1}{2}^+)$

Mass  $m = 1115.683 \pm 0.006$  MeV  
( $m_\Lambda - m_\pi$ ) /  $m_\Lambda = (-0.1 \pm 1.1) \times 10^{-5}$  ( $S = 1.6$ )  
Mean life  $\tau = (2.632 \pm 0.020) \times 10^{-10}$  s ( $S = 1.6$ )  
( $\tau_\Lambda - \tau_{\bar{\Lambda}}$ ) /  $\tau_\Lambda = -0.001 \pm 0.009$   
 $c\tau = 7.89$  cm  
Magnetic moment  $\mu = -0.613 \pm 0.004 \mu_N$   
Electric dipole moment  $d < 1.5 \times 10^{-16}$  e cm, CL = 95%

Decay parameters

$p\pi^-$	$\alpha_- = 0.642 \pm 0.013$
$\bar{p}\pi^+$	$\alpha_+ = -0.71 \pm 0.08$
$p\pi^-$	$\phi_- = (-6.5 \pm 3.5)^\circ$
"	$\gamma_- = 0.76$ [n]
"	$\Delta_- = (8 \pm 4)^\circ$ [n]
$n\pi^0$	$\alpha_0 = 0.65 \pm 0.04$
$p e^- \bar{\nu}_e$	$g_A/g_V = -0.718 \pm 0.015$ [l]

$\Lambda$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$p\pi^-$	(63.9 $\pm$ 0.5 ) %		101
$n\pi^0$	(35.8 $\pm$ 0.5 ) %		104
$n\gamma$	( 1.75 $\pm$ 0.15 ) $\times 10^{-3}$		162
$p\pi^- \gamma$	[o] ( 8.4 $\pm$ 1.4 ) $\times 10^{-4}$		101
$p e^- \bar{\nu}_e$	( 8.32 $\pm$ 0.14 ) $\times 10^{-4}$		163
$p\mu^- \bar{\nu}_\mu$	( 1.57 $\pm$ 0.35 ) $\times 10^{-4}$		131

Lepton ( $L$ ) and/or Baryon ( $B$ ) number violating decay modes

$\pi^+ e^-$	$L, B$	< 6	$\times 10^{-7}$	90%	549
$\pi^+ \mu^-$	$L, B$	< 6	$\times 10^{-7}$	90%	544
$\pi^- e^+$	$L, B$	< 4	$\times 10^{-7}$	90%	549
$\pi^- \mu^+$	$L, B$	< 6	$\times 10^{-7}$	90%	544
$K^+ e^-$	$L, B$	< 2	$\times 10^{-6}$	90%	449
$K^+ \mu^-$	$L, B$	< 3	$\times 10^{-6}$	90%	441
$K^- e^+$	$L, B$	< 2	$\times 10^{-6}$	90%	449
$K^- \mu^+$	$L, B$	< 3	$\times 10^{-6}$	90%	441
$K_S^0 \nu$	$L, B$	< 2	$\times 10^{-5}$	90%	447
$\bar{p}\pi^+$	$B$	< 9	$\times 10^{-7}$	90%	101

$\Lambda(1405) 1/2^-$

$I(J^P) = 0(\frac{1}{2}^-)$

Mass  $m = 1405.1^{+1.3}_{-1.0}$  MeV  
Full width  $\Gamma = 50.5 \pm 2.0$  MeV  
Below  $\bar{K}N$  threshold

$\Lambda(1405)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Sigma\pi$	100 %	155

$\Lambda(1520) 3/2^-$

$I(J^P) = 0(\frac{3}{2}^-)$

Mass  $m = 1519.5 \pm 1.0$  MeV [p]  
Full width  $\Gamma = 15.6 \pm 1.0$  MeV [p]

$\Lambda(1520)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	(45 $\pm$ 1 ) %	243
$\Sigma\pi$	(42 $\pm$ 1 ) %	268
$\Lambda\pi\pi$	(10 $\pm$ 1 ) %	259
$\Sigma\pi\pi$	( 0.9 $\pm$ 0.1 ) %	169
$\Lambda\gamma$	( 0.85 $\pm$ 0.15 ) %	350

$\Lambda(1600) 1/2^+$

$I(J^P) = 0(\frac{1}{2}^+)$

Mass  $m = 1560$  to 1700 ( $\approx$  1600) MeV  
Full width  $\Gamma = 50$  to 250 ( $\approx$  150) MeV

## Baryon Summary Table

$\Lambda(1600)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	15–30 %	343
$\Sigma\pi$	10–60 %	338

$$\Lambda(1670) \ 1/2^- \quad I(J^P) = 0(\frac{1}{2}^-)$$

Mass  $m = 1660$  to  $1680$  ( $\approx 1670$ ) MeV  
Full width  $\Gamma = 25$  to  $50$  ( $\approx 35$ ) MeV

$\Lambda(1670)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	20–30 %	414
$\Sigma\pi$	25–55 %	394
$\Lambda\eta$	10–25 %	69
$N\bar{K}^*(892)$ , $S=3/2$ , $D$ -wave	(5 $\pm$ 4) %	†

$$\Lambda(1690) \ 3/2^- \quad I(J^P) = 0(\frac{3}{2}^-)$$

Mass  $m = 1685$  to  $1695$  ( $\approx 1690$ ) MeV  
Full width  $\Gamma = 50$  to  $70$  ( $\approx 60$ ) MeV

$\Lambda(1690)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	20–30 %	433
$\Sigma\pi$	20–40 %	410
$\Lambda\pi\pi$	$\sim 25$ %	419
$\Sigma\pi\pi$	$\sim 20$ %	358

$$\Lambda(1800) \ 1/2^- \quad I(J^P) = 0(\frac{1}{2}^-)$$

Mass  $m = 1720$  to  $1850$  ( $\approx 1800$ ) MeV  
Full width  $\Gamma = 200$  to  $400$  ( $\approx 300$ ) MeV

$\Lambda(1800)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	25–40 %	528
$\Sigma\pi$	seen	494
$\Sigma(1385)\pi$	seen	349
$\Lambda\eta$	(6 $\pm$ 5) %	326
$N\bar{K}^*(892)$	seen	†

$$\Lambda(1810) \ 1/2^+ \quad I(J^P) = 0(\frac{1}{2}^+)$$

Mass  $m = 1750$  to  $1850$  ( $\approx 1810$ ) MeV  
Full width  $\Gamma = 50$  to  $250$  ( $\approx 150$ ) MeV

$\Lambda(1810)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	20–50 %	537
$\Sigma\pi$	10–40 %	501
$\Sigma(1385)\pi$	seen	357
$N\bar{K}^*(892)$	30–60 %	†

$$\Lambda(1820) \ 5/2^+ \quad I(J^P) = 0(\frac{5}{2}^+)$$

Mass  $m = 1815$  to  $1825$  ( $\approx 1820$ ) MeV  
Full width  $\Gamma = 70$  to  $90$  ( $\approx 80$ ) MeV

$\Lambda(1820)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	55–65 %	545
$\Sigma\pi$	8–14 %	509
$\Sigma(1385)\pi$	5–10 %	366
$N\bar{K}^*(892)$ , $S=3/2$ , $P$ -wave	(3.0 $\pm$ 1.0) %	†

$$\Lambda(1830) \ 5/2^- \quad I(J^P) = 0(\frac{5}{2}^-)$$

Mass  $m = 1810$  to  $1830$  ( $\approx 1830$ ) MeV  
Full width  $\Gamma = 60$  to  $110$  ( $\approx 95$ ) MeV

$\Lambda(1830)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	3–10 %	553
$\Sigma\pi$	35–75 %	516
$\Sigma(1385)\pi$	>15 %	374
$\Sigma(1385)\pi$ , $D$ -wave	(52 $\pm$ 6) %	374

$$\Lambda(1890) \ 3/2^+ \quad I(J^P) = 0(\frac{3}{2}^+)$$

Mass  $m = 1850$  to  $1910$  ( $\approx 1890$ ) MeV  
Full width  $\Gamma = 60$  to  $200$  ( $\approx 100$ ) MeV

$\Lambda(1890)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	20–35 %	599
$\Sigma\pi$	3–10 %	560
$\Sigma(1385)\pi$	seen	423
$N\bar{K}^*(892)$	seen	236

$$\Lambda(2100) \ 7/2^- \quad I(J^P) = 0(\frac{7}{2}^-)$$

Mass  $m = 2090$  to  $2110$  ( $\approx 2100$ ) MeV  
Full width  $\Gamma = 100$  to  $250$  ( $\approx 200$ ) MeV

$\Lambda(2100)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	25–35 %	751
$\Sigma\pi$	$\sim 5$ %	705
$\Lambda\eta$	<3 %	617
$\Xi K$	<3 %	491
$\Lambda\omega$	<8 %	443
$N\bar{K}^*(892)$	10–20 %	515

$$\Lambda(2110) \ 5/2^+ \quad I(J^P) = 0(\frac{5}{2}^+)$$

Mass  $m = 2090$  to  $2140$  ( $\approx 2110$ ) MeV  
Full width  $\Gamma = 150$  to  $250$  ( $\approx 200$ ) MeV

$\Lambda(2110)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	5–25 %	757
$\Sigma\pi$	10–40 %	711
$\Lambda\omega$	seen	455
$\Sigma(1385)\pi$	seen	591
$N\bar{K}^*(892)$	10–60 %	525

$$\Lambda(2350) \ 9/2^+ \quad I(J^P) = 0(\frac{9}{2}^+)$$

Mass  $m = 2340$  to  $2370$  ( $\approx 2350$ ) MeV  
Full width  $\Gamma = 100$  to  $250$  ( $\approx 150$ ) MeV

$\Lambda(2350)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$N\bar{K}$	$\sim 12$ %	915
$\Sigma\pi$	$\sim 10$ %	867

## $\Sigma$ BARYONS ( $S = -1$ , $I = 1$ )

$$\Sigma^+ = u u s, \quad \Sigma^0 = u d s, \quad \Sigma^- = d d s$$

$$\Sigma^+ \quad I(J^P) = 1(\frac{1}{2}^+)$$

Mass  $m = 1189.37 \pm 0.07$  MeV ( $S = 2.2$ )

Mean life  $\tau = (0.8018 \pm 0.0026) \times 10^{-10}$  s

$$c\tau = 2.404 \text{ cm}$$

$$(\tau_{\Sigma^+} - \tau_{\Sigma^-}) / \tau_{\Sigma^+} = -0.0006 \pm 0.0012$$

Magnetic moment  $\mu = 2.458 \pm 0.010 \mu_N$  ( $S = 2.1$ )

$$(\mu_{\Sigma^+} + \mu_{\Sigma^-}) / \mu_{\Sigma^+} = 0.014 \pm 0.015$$

$$\Gamma(\Sigma^+ \rightarrow n \ell^+ \nu) / \Gamma(\Sigma^- \rightarrow n \ell^- \bar{\nu}) < 0.043$$

# Baryon Summary Table

## Decay parameters

$p\pi^0$	$\alpha_0 = -0.980^{+0.017}_{-0.015}$
"	$\phi_0 = (36 \pm 34)^\circ$
"	$\gamma_0 = 0.16 [n]$
"	$\Delta_0 = (187 \pm 6)^\circ [n]$
$n\pi^+$	$\alpha_+ = 0.068 \pm 0.013$
"	$\phi_+ = (167 \pm 20)^\circ \quad (S = 1.1)$
"	$\gamma_+ = -0.97 [n]$
"	$\Delta_+ = (-73^{+133}_{-10})^\circ [n]$
$p\gamma$	$\alpha_\gamma = -0.76 \pm 0.08$

$\Sigma^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$p\pi^0$	$(51.57 \pm 0.30) \%$		189
$n\pi^+$	$(48.31 \pm 0.30) \%$		185
$p\gamma$	$(1.23 \pm 0.05) \times 10^{-3}$		225
$n\pi^+\gamma$	$[q] \quad (4.5 \pm 0.5) \times 10^{-4}$		185
$\Lambda e^+ \nu_e$	$(2.0 \pm 0.5) \times 10^{-5}$		71

## $\Delta S = \Delta Q$ (SQ) violating modes or $\Delta S = 1$ weak neutral current (S1) modes

$n e^+ \nu_e$	SQ	$< 5 \times 10^{-6}$	90%	224
$n\mu^+ \nu_\mu$	SQ	$< 3.0 \times 10^{-5}$	90%	202
$p e^+ e^-$	S1	$< 7 \times 10^{-6}$		225
$p\mu^+ \mu^-$	S1	$(9^{+9}_{-8}) \times 10^{-8}$		121

$$\Sigma^0 \quad I(J^P) = 1(\frac{1}{2}^+)$$

Mass  $m = 1192.642 \pm 0.024$  MeV  
 $m_{\Sigma^-} - m_{\Sigma^0} = 4.807 \pm 0.035$  MeV  $(S = 1.1)$   
 $m_{\Sigma^0} - m_\Lambda = 76.959 \pm 0.023$  MeV  
Mean life  $\tau = (7.4 \pm 0.7) \times 10^{-20}$  s  
 $c\tau = 2.22 \times 10^{-11}$  m  
Transition magnetic moment  $|\mu_{\Sigma\Lambda}| = 1.61 \pm 0.08 \mu_N$

$\Sigma^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\Lambda\gamma$	100 %		74
$\Lambda\gamma\gamma$	$< 3 \%$	90%	74
$\Lambda e^+ e^-$	$[q] \quad 5 \times 10^{-3}$		74

$$\Sigma^- \quad I(J^P) = 1(\frac{1}{2}^+)$$

Mass  $m = 1197.449 \pm 0.030$  MeV  $(S = 1.2)$   
 $m_{\Sigma^-} - m_{\Sigma^+} = 8.08 \pm 0.08$  MeV  $(S = 1.9)$   
 $m_{\Sigma^-} - m_\Lambda = 81.766 \pm 0.030$  MeV  $(S = 1.2)$   
Mean life  $\tau = (1.479 \pm 0.011) \times 10^{-10}$  s  $(S = 1.3)$   
 $c\tau = 4.434$  cm  
Magnetic moment  $\mu = -1.160 \pm 0.025 \mu_N \quad (S = 1.7)$   
 $\Sigma^-$  charge radius =  $0.78 \pm 0.10$  fm

## Decay parameters

$n\pi^-$	$\alpha_- = -0.068 \pm 0.008$
"	$\phi_- = (10 \pm 15)^\circ$
"	$\gamma_- = 0.98 [n]$
"	$\Delta_- = (249^{+12}_{-120})^\circ [n]$
$n e^- \bar{\nu}_e$	$g_A/g_V = 0.340 \pm 0.017 [l]$
"	$f_2(0)/f_1(0) = 0.97 \pm 0.14$
"	$D = 0.11 \pm 0.10$
$\Lambda e^- \bar{\nu}_e$	$g_V/g_A = 0.01 \pm 0.10 [l] \quad (S = 1.5)$
"	$g_{WM}/g_A = 2.4 \pm 1.7 [l]$

## $\Sigma^-$ DECAY MODES

	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$n\pi^-$	$(99.848 \pm 0.005) \%$	193
$n\pi^-\gamma$	$[q] \quad (4.6 \pm 0.6) \times 10^{-4}$	193
$n e^- \bar{\nu}_e$	$(1.017 \pm 0.034) \times 10^{-3}$	230
$n\mu^- \bar{\nu}_\mu$	$(4.5 \pm 0.4) \times 10^{-4}$	210
$\Lambda e^- \bar{\nu}_e$	$(5.73 \pm 0.27) \times 10^{-5}$	79

## $\Sigma(1385) 3/2^+$

$$I(J^P) = 1(\frac{3}{2}^+)$$

$\Sigma(1385)^+$  mass  $m = 1382.80 \pm 0.35$  MeV  $(S = 1.9)$   
 $\Sigma(1385)^0$  mass  $m = 1383.7 \pm 1.0$  MeV  $(S = 1.4)$   
 $\Sigma(1385)^-$  mass  $m = 1387.2 \pm 0.5$  MeV  $(S = 2.2)$   
 $\Sigma(1385)^+$  full width  $\Gamma = 36.0 \pm 0.7$  MeV  
 $\Sigma(1385)^0$  full width  $\Gamma = 36 \pm 5$  MeV  
 $\Sigma(1385)^-$  full width  $\Gamma = 39.4 \pm 2.1$  MeV  $(S = 1.7)$   
Below  $\bar{K}N$  threshold

$\Sigma(1385)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\Lambda\pi$	$(87.0 \pm 1.5) \%$		208
$\Sigma\pi$	$(11.7 \pm 1.5) \%$		129
$\Lambda\gamma$	$(1.25^{+0.13}_{-0.12}) \%$		241
$\Sigma^+\gamma$	$(7.0 \pm 1.7) \times 10^{-3}$		180
$\Sigma^-\gamma$	$< 2.4 \times 10^{-4}$	90%	173

## $\Sigma(1660) 1/2^+$

$$I(J^P) = 1(\frac{1}{2}^+)$$

Mass  $m = 1630$  to  $1690$  ( $\approx 1660$ ) MeV  
Full width  $\Gamma = 40$  to  $200$  ( $\approx 100$ ) MeV

$\Sigma(1660)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	10–30 %	405
$\Lambda\pi$	seen	440
$\Sigma\pi$	seen	387

## $\Sigma(1670) 3/2^-$

$$I(J^P) = 1(\frac{3}{2}^-)$$

Mass  $m = 1665$  to  $1685$  ( $\approx 1670$ ) MeV  
Full width  $\Gamma = 40$  to  $80$  ( $\approx 60$ ) MeV

$\Sigma(1670)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	7–13 %	414
$\Lambda\pi$	5–15 %	448
$\Sigma\pi$	30–60 %	394

## $\Sigma(1750) 1/2^-$

$$I(J^P) = 1(\frac{1}{2}^-)$$

Mass  $m = 1730$  to  $1800$  ( $\approx 1750$ ) MeV  
Full width  $\Gamma = 60$  to  $160$  ( $\approx 90$ ) MeV

$\Sigma(1750)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	10–40 %	486
$\Lambda\pi$	seen	507
$\Sigma\pi$	$< 8 \%$	456
$\Sigma\eta$	15–55 %	98
$N\bar{K}^*(892), S=1/2$	$(8 \pm 4) \%$	†

## $\Sigma(1775) 5/2^-$

$$I(J^P) = 1(\frac{5}{2}^-)$$

Mass  $m = 1770$  to  $1780$  ( $\approx 1775$ ) MeV  
Full width  $\Gamma = 105$  to  $135$  ( $\approx 120$ ) MeV

## Baryon Summary Table

$\Sigma(1775)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	37–43%	508
$\Lambda\pi$	14–20%	525
$\Sigma\pi$	2–5%	475
$\Sigma(1385)\pi$	8–12%	327
$\Lambda(1520)\pi$ , $P$ -wave	17–23%	201

 **$\Sigma(1915)$   $5/2^+$** 

$$I(J^P) = 1(\frac{5}{2}^+)$$

Mass  $m = 1900$  to  $1935$  ( $\approx 1915$ ) MeVFull width  $\Gamma = 80$  to  $160$  ( $\approx 120$ ) MeV

$\Sigma(1915)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	5–15 %	618
$\Lambda\pi$	seen	623
$\Sigma\pi$	seen	577
$\Sigma(1385)\pi$	<5 %	443

 **$\Sigma(1940)$   $3/2^-$** 

$$I(J^P) = 1(\frac{3}{2}^-)$$

Mass  $m = 1900$  to  $1950$  ( $\approx 1940$ ) MeVFull width  $\Gamma = 150$  to  $300$  ( $\approx 220$ ) MeV

$\Sigma(1940)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	<20 %	637
$\Lambda\pi$	seen	640
$\Sigma\pi$	seen	595
$\Sigma(1385)\pi$	seen	463
$\Lambda(1520)\pi$	seen	355
$\Delta(1232)\bar{K}$	seen	410
$N\bar{K}^*(892)$	seen	322

 **$\Sigma(2030)$   $7/2^+$** 

$$I(J^P) = 1(\frac{7}{2}^+)$$

Mass  $m = 2025$  to  $2040$  ( $\approx 2030$ ) MeVFull width  $\Gamma = 150$  to  $200$  ( $\approx 180$ ) MeV

$\Sigma(2030)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	17–23 %	702
$\Lambda\pi$	17–23 %	700
$\Sigma\pi$	5–10 %	657
$\Xi K$	<2 %	422
$\Sigma(1385)\pi$	5–15 %	532
$\Lambda(1520)\pi$	10–20 %	430
$\Delta(1232)\bar{K}$	10–20 %	498
$N\bar{K}^*(892)$	<5 %	439

 **$\Sigma(2250)$** 

$$I(J^P) = 1(?^?)$$

Mass  $m = 2210$  to  $2280$  ( $\approx 2250$ ) MeVFull width  $\Gamma = 60$  to  $150$  ( $\approx 100$ ) MeV

$\Sigma(2250)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$N\bar{K}$	<10 %	851
$\Lambda\pi$	seen	842
$\Sigma\pi$	seen	803

 **$\Xi$  BARYONS**  
 **$(S = -2, I = 1/2)$** 

$$\Xi^0 = u s s, \quad \Xi^- = d s s$$

 **$\Xi^0$** 

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

 $P$  is not yet measured; + is the quark model prediction.Mass  $m = 1314.86 \pm 0.20$  MeV $m_{\Xi^-} - m_{\Xi^0} = 6.85 \pm 0.21$  MeVMean life  $\tau = (2.90 \pm 0.09) \times 10^{-10}$  s $c\tau = 8.71$  cmMagnetic moment  $\mu = -1.250 \pm 0.014 \mu_N$ **Decay parameters**

$$\Lambda\pi^0 \quad \alpha = -0.406 \pm 0.013$$

$$" \quad \phi = (21 \pm 12)^\circ$$

$$" \quad \gamma = 0.85 [n]$$

$$" \quad \Delta = (218^{+12}_{-19})^\circ [n]$$

$$\Lambda\gamma \quad \alpha = -0.70 \pm 0.07$$

$$\Lambda e^+ e^- \quad \alpha = -0.8 \pm 0.2$$

$$\Sigma^0 \gamma \quad \alpha = -0.69 \pm 0.06$$

$$\Sigma^+ e^- \bar{\nu}_e \quad g_1(0)/f_1(0) = 1.22 \pm 0.05$$

$$\Sigma^+ e^- \bar{\nu}_e \quad f_2(0)/f_1(0) = 2.0 \pm 0.9$$

 **$\Xi^0$  DECAY MODES**

	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\Lambda\pi^0$	$(99.524 \pm 0.012) \%$		135
$\Lambda\gamma$	$(1.17 \pm 0.07) \times 10^{-3}$		184
$\Lambda e^+ e^-$	$(7.6 \pm 0.6) \times 10^{-6}$		184
$\Sigma^0 \gamma$	$(3.33 \pm 0.10) \times 10^{-3}$		117
$\Sigma^+ e^- \bar{\nu}_e$	$(2.52 \pm 0.08) \times 10^{-4}$		120
$\Sigma^+ \mu^- \bar{\nu}_\mu$	$(2.33 \pm 0.35) \times 10^{-6}$		64

 **$\Delta S = \Delta Q$  (SQ) violating modes or  
 $\Delta S = 2$  forbidden (S2) modes**

$\Sigma^- e^+ \nu_e$	SQ	< 9	$\times 10^{-4}$	90%	112
$\Sigma^- \mu^+ \nu_\mu$	SQ	< 9	$\times 10^{-4}$	90%	49
$p\pi^-$	S2	< 8	$\times 10^{-6}$	90%	299
$p e^- \bar{\nu}_e$	S2	< 1.3	$\times 10^{-3}$		323
$p\mu^- \bar{\nu}_\mu$	S2	< 1.3	$\times 10^{-3}$		309

 **$\Xi^-$** 

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

 $P$  is not yet measured; + is the quark model prediction.Mass  $m = 1321.71 \pm 0.07$  MeV $(m_{\Xi^-} - m_{\Xi^+}) / m_{\Xi^-} = (-3 \pm 9) \times 10^{-5}$ Mean life  $\tau = (1.639 \pm 0.015) \times 10^{-10}$  s $c\tau = 4.91$  cm $(\tau_{\Xi^-} - \tau_{\Xi^+}) / \tau_{\Xi^-} = -0.01 \pm 0.07$ Magnetic moment  $\mu = -0.6507 \pm 0.0025 \mu_N$  $(\mu_{\Xi^-} + \mu_{\Xi^+}) / |\mu_{\Xi^-}| = +0.01 \pm 0.05$ **Decay parameters**

$$\Lambda\pi^- \quad \alpha = -0.458 \pm 0.012 \quad (S = 1.8)$$

$$[\alpha(\Xi^-)\alpha_-(\Lambda) - \alpha(\Xi^+)\alpha_+(\bar{\Lambda})] / [\text{sum}] = (0 \pm 7) \times 10^{-4}$$

$$" \quad \phi = (-2.1 \pm 0.8)^\circ$$

$$" \quad \gamma = 0.89 [n]$$

$$" \quad \Delta = (175.9 \pm 1.5)^\circ [n]$$

$$\Lambda e^- \bar{\nu}_e \quad g_A/g_V = -0.25 \pm 0.05 [l]$$

# Baryon Summary Table

$\Xi^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\Lambda\pi^-$	(99.887±0.035) %		140
$\Sigma^-\gamma$	( 1.27 ±0.23 ) × 10 <sup>-4</sup>		118
$\Lambda e^-\bar{\nu}_e$	( 5.63 ±0.31 ) × 10 <sup>-4</sup>		190
$\Lambda\mu^-\bar{\nu}_\mu$	( 3.5 <sup>+3.5</sup> <sub>-2.2</sub> ) × 10 <sup>-4</sup>		163
$\Sigma^0 e^-\bar{\nu}_e$	( 8.7 ±1.7 ) × 10 <sup>-5</sup>		123
$\Sigma^0\mu^-\bar{\nu}_\mu$	< 8 × 10 <sup>-4</sup>	90%	70
$\Xi^0 e^-\bar{\nu}_e$	< 2.3 × 10 <sup>-3</sup>	90%	7
<b><math>\Delta S = 2</math> forbidden (<math>S2</math>) modes</b>			
$n\pi^-$	$S2$ < 1.9 × 10 <sup>-5</sup>	90%	304
$n e^-\bar{\nu}_e$	$S2$ < 3.2 × 10 <sup>-3</sup>	90%	327
$n\mu^-\bar{\nu}_\mu$	$S2$ < 1.5 %	90%	314
$p\pi^-\pi^-$	$S2$ < 4 × 10 <sup>-4</sup>	90%	223
$p\pi^- e^-\bar{\nu}_e$	$S2$ < 4 × 10 <sup>-4</sup>	90%	305
$p\pi^-\mu^-\bar{\nu}_\mu$	$S2$ < 4 × 10 <sup>-4</sup>	90%	251
$p\mu^-\mu^-$	$L$ < 4 × 10 <sup>-8</sup>	90%	272

$\Xi(1530) \ 3/2^+$

$I(J^P) = \frac{1}{2}(\frac{3}{2}^+)$

$\Xi(1530)^0$  mass  $m = 1531.80 \pm 0.32$  MeV ( $S = 1.3$ )  
 $\Xi(1530)^-$  mass  $m = 1535.0 \pm 0.6$  MeV  
 $\Xi(1530)^0$  full width  $\Gamma = 9.1 \pm 0.5$  MeV  
 $\Xi(1530)^-$  full width  $\Gamma = 9.9^{+1.7}_{-1.9}$  MeV

$\Xi(1530)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\Xi\pi$	100 %		158
$\Xi\gamma$	<4 %	90%	202

$\Xi(1690)$

$I(J^P) = \frac{1}{2}(?^?)$

Mass  $m = 1690 \pm 10$  MeV [ $p$ ]  
Full width  $\Gamma < 30$  MeV

$\Xi(1690)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda\bar{K}$	seen	240
$\Sigma\bar{K}$	seen	70
$\Xi\pi$	seen	311
$\Xi^-\pi^+\pi^-$	possibly seen	213

$\Xi(1820) \ 3/2^-$

$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$

Mass  $m = 1823 \pm 5$  MeV [ $p$ ]  
Full width  $\Gamma = 24^{+15}_{-10}$  MeV [ $p$ ]

$\Xi(1820)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda\bar{K}$	large	402
$\Sigma\bar{K}$	small	324
$\Xi\pi$	small	421
$\Xi(1530)\pi$	small	237

$\Xi(1950)$

$I(J^P) = \frac{1}{2}(?^?)$

Mass  $m = 1950 \pm 15$  MeV [ $p$ ]  
Full width  $\Gamma = 60 \pm 20$  MeV [ $p$ ]

$\Xi(1950)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda\bar{K}$	seen	522
$\Sigma\bar{K}$	possibly seen	460
$\Xi\pi$	seen	519

$\Xi(2030)$

$I(J^P) = \frac{1}{2}(\geq \frac{5}{2}^?)$

Mass  $m = 2025 \pm 5$  MeV [ $p$ ]  
Full width  $\Gamma = 20^{+15}_{-5}$  MeV [ $p$ ]

$\Xi(2030)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda\bar{K}$	~ 20 %	585
$\Sigma\bar{K}$	~ 80 %	529
$\Xi\pi$	small	574
$\Xi(1530)\pi$	small	416
$\Lambda\bar{K}\pi$	small	499
$\Sigma\bar{K}\pi$	small	428

$\Omega$  BARYONS  
( $S = -3, I = 0$ )  
 $\Omega^- = sss$

$\Omega^-$

$I(J^P) = 0(\frac{3}{2}^+)$

$J^P = \frac{3}{2}^+$  is the quark-model prediction; and  $J = 3/2$  is fairly well established.  
Mass  $m = 1672.45 \pm 0.29$  MeV  
 $(m_{\Omega^-} - m_{\bar{\Omega}^+}) / m_{\Omega^-} = (-1 \pm 8) \times 10^{-5}$   
Mean life  $\tau = (0.821 \pm 0.011) \times 10^{-10}$  s  
 $c\tau = 2.461$  cm  
 $(\tau_{\Omega^-} - \tau_{\bar{\Omega}^+}) / \tau_{\Omega^-} = 0.00 \pm 0.05$   
Magnetic moment  $\mu = -2.02 \pm 0.05 \mu_N$

Decay parameters	
$\Lambda K^-$	$\alpha = 0.0180 \pm 0.0024$
$\Lambda K^-, \bar{\Lambda} K^+$	$(\alpha + \bar{\alpha}) / (\alpha - \bar{\alpha}) = -0.02 \pm 0.13$
$\Xi^0 \pi^-$	$\alpha = 0.09 \pm 0.14$
$\Xi^- \pi^0$	$\alpha = 0.05 \pm 0.21$

$\Omega^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$p$ (MeV/c)
$\Lambda K^-$	(67.8±0.7) %		211
$\Xi^0 \pi^-$	(23.6±0.7) %		294
$\Xi^- \pi^0$	( 8.6±0.4 ) %		289
$\Xi^- \pi^+ \pi^-$	( 3.7 <sup>+0.7</sup> <sub>-0.6</sub> ) × 10 <sup>-4</sup>		189
$\Xi(1530)^0 \pi^-$	< 7 × 10 <sup>-5</sup>	90%	17
$\Xi^0 e^-\bar{\nu}_e$	( 5.6±2.8 ) × 10 <sup>-3</sup>		319
$\Xi^- \gamma$	< 4.6 × 10 <sup>-4</sup>	90%	314

<b><math>\Delta S = 2</math> forbidden (<math>S2</math>) modes</b>			
$\Lambda\pi^-$	$S2$ < 2.9 × 10 <sup>-6</sup>	90%	449

$\Omega(2250)^-$

$I(J^P) = 0(?^?)$

Mass  $m = 2252 \pm 9$  MeV  
Full width  $\Gamma = 55 \pm 18$  MeV

$\Omega(2250)^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Xi^- \pi^+ K^-$	seen	532
$\Xi(1530)^0 K^-$	seen	437

CHARMED BARYONS  
( $C = +1$ )  
 $\Lambda_c^+ = udc, \ \Sigma_c^{++} = uuc, \ \Sigma_c^+ = udc, \ \Sigma_c^0 = ddc,$   
 $\Xi_c^+ = usc, \ \Xi_c^0 = dsc, \ \Omega_c^0 = ssc$

$\Lambda_c^+$

$I(J^P) = 0(\frac{1}{2}^+)$

$J$  is not well measured;  $\frac{1}{2}$  is the quark-model prediction.  
Mass  $m = 2286.46 \pm 0.14$  MeV  
Mean life  $\tau = (200 \pm 6) \times 10^{-15}$  s ( $S = 1.6$ )  
 $c\tau = 59.9 \mu\text{m}$



## Baryon Summary Table

## Decay asymmetry parameters

$\Lambda\pi^+$	$\alpha = -0.91 \pm 0.15$
$\Sigma^+\pi^0$	$\alpha = -0.45 \pm 0.32$
$\Lambda\ell^+\nu_\ell$	$\alpha = -0.86 \pm 0.04$
$(\alpha + \bar{\alpha})/(\alpha - \bar{\alpha})$ in $\Lambda_C^+ \rightarrow \Lambda\pi^+, \bar{\Lambda}_C^- \rightarrow \bar{\Lambda}\pi^-$	$= -0.07 \pm 0.31$
$(\alpha + \bar{\alpha})/(\alpha - \bar{\alpha})$ in $\Lambda_C^+ \rightarrow \Lambda e^+\nu_e, \bar{\Lambda}_C^- \rightarrow \bar{\Lambda}e^-\bar{\nu}_e$	$= 0.00 \pm 0.04$

$\Lambda_C^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$p$ (MeV/c)
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Hadronic modes with a  $p$ :  $S = -1$  final states

$pK_S^0$	( 1.58 ± 0.08 ) %	S=1.2	873
$pK^-\pi^+$	( 6.35 ± 0.33 ) %	S=1.4	823
$p\bar{K}^*(892)^0$	[r] ( 1.98 ± 0.28 ) %		685
$\Delta(1232)^{++}K^-$	( 1.09 ± 0.25 ) %		710
$\Lambda(1520)\pi^+$	[r] ( 2.2 ± 0.5 ) %		627
$pK^-\pi^+$ nonresonant	( 3.5 ± 0.4 ) %		823
$pK_S^0\pi^0$	( 1.99 ± 0.13 ) %	S=1.1	823
$p\bar{K}^0\eta$	( 1.6 ± 0.4 ) %		568
$pK_S^0\pi^+\pi^-$	( 1.66 ± 0.12 ) %	S=1.1	754
$pK^-\pi^+\pi^0$	( 4.9 ± 0.4 ) %	S=1.3	759
$pK^*(892)^-\pi^+$	[r] ( 1.5 ± 0.5 ) %		580
$p(K^-\pi^+)_{\text{nonresonant}}\pi^0$	( 4.6 ± 0.9 ) %		759
$\Delta(1232)\bar{K}^*(892)$	seen		419
$pK^-\pi^+\pi^0$	( 1.4 ± 1.0 ) × 10 <sup>-3</sup>		671
$pK^-\pi^+2\pi^0$	( 1.0 ± 0.5 ) %		678

Hadronic modes with a  $p$ :  $S = 0$  final states

$p\pi^+\pi^-$	( 4.4 ± 2.3 ) × 10 <sup>-3</sup>		927
$p f_0(980)$	[r] ( 3.5 ± 2.3 ) × 10 <sup>-3</sup>		614
$p2\pi^+2\pi^-$	( 2.3 ± 1.5 ) × 10 <sup>-3</sup>		852
$pK^+K^-$	( 10 ± 4 ) × 10 <sup>-4</sup>		616
$p\phi$	[r] ( 1.04 ± 0.21 ) × 10 <sup>-3</sup>		590
$pK^+K^-$ non- $\phi$	( 4.4 ± 1.8 ) × 10 <sup>-4</sup>		616

Hadronic modes with a hyperon:  $S = -1$  final states

$\Lambda\pi^+$	( 1.30 ± 0.07 ) %	S=1.2	864
$\Lambda\pi^+\pi^0$	( 7.1 ± 0.4 ) %	S=1.2	844
$\Lambda\rho^+$	< 6 %	CL=95%	636
$\Lambda\pi^-2\pi^+$	( 3.7 ± 0.4 ) %	S=1.9	807
$\Sigma(1385)^+\pi^+\pi^-, \Sigma^{*+} \rightarrow \Lambda\pi^+$	( 1.0 ± 0.5 ) %		688
$\Sigma(1385)^-2\pi^+, \Sigma^{*-} \rightarrow \Lambda\pi^-$	( 7.8 ± 1.6 ) × 10 <sup>-3</sup>		688
$\Lambda\pi^+\rho^0$	( 1.5 ± 0.6 ) %		524
$\Sigma(1385)^+\rho^0, \Sigma^{*+} \rightarrow \Lambda\pi^+$	( 5 ± 4 ) × 10 <sup>-3</sup>		363
$\Lambda\pi^-2\pi^+$ nonresonant	< 1.1 %	CL=90%	807
$\Lambda\pi^-\pi^02\pi^+$ total	( 2.3 ± 0.8 ) %		757
$\Lambda\pi^+\eta$	[r] ( 2.3 ± 0.5 ) %		691
$\Sigma(1385)^+\eta$	[r] ( 1.08 ± 0.32 ) %		570
$\Lambda\pi^+\omega$	[r] ( 1.5 ± 0.5 ) %		517
$\Lambda\pi^-\pi^02\pi^+, \text{ no } \eta \text{ or } \omega$	< 8 × 10 <sup>-3</sup>	CL=90%	757
$\Lambda K^+\bar{K}^0$	( 5.7 ± 1.1 ) × 10 <sup>-3</sup>	S=2.0	443
$\Xi(1690)^0K^+, \Xi^{*0} \rightarrow \Lambda\bar{K}^0$	( 1.6 ± 0.5 ) × 10 <sup>-3</sup>		286
$\Sigma^0\pi^+$	( 1.29 ± 0.07 ) %	S=1.1	825
$\Sigma^+\pi^0$	( 1.24 ± 0.10 ) %		827
$\Sigma^+\eta$	( 7.0 ± 2.3 ) × 10 <sup>-3</sup>		713
$\Sigma^+\pi^+\pi^-$	( 4.57 ± 0.29 ) %	S=1.2	804
$\Sigma^+\rho^0$	< 1.7 %	CL=95%	575
$\Sigma^-2\pi^+$	( 2.1 ± 0.4 ) %		799
$\Sigma^0\pi^+\pi^0$	( 2.3 ± 0.9 ) %		803
$\Sigma^0\pi^-2\pi^+$	( 1.13 ± 0.29 ) %		763
$\Sigma^+\pi^+\pi^-\pi^0$	—		767
$\Sigma^+\omega$	[r] ( 1.74 ± 0.21 ) %		569
$\Sigma^+K^+K^-$	( 3.6 ± 0.4 ) × 10 <sup>-3</sup>		349
$\Sigma^+\phi$	[r] ( 4.0 ± 0.6 ) × 10 <sup>-3</sup>	S=1.1	295
$\Xi(1690)^0K^+, \Xi^{*0} \rightarrow \Sigma^+K^-$	( 1.03 ± 0.26 ) × 10 <sup>-3</sup>		286
$\Sigma^+K^+K^-$ nonresonant	< 8 × 10 <sup>-4</sup>	CL=90%	349
$\Xi^0K^+$	( 5.0 ± 1.2 ) × 10 <sup>-3</sup>		653
$\Xi^-K^+\pi^+$	( 6.2 ± 0.6 ) × 10 <sup>-3</sup>	S=1.1	565
$\Xi(1530)^0K^+$	[r] ( 3.3 ± 0.9 ) × 10 <sup>-3</sup>		473

Hadronic modes with a hyperon:  $S = 0$  final states

$\Lambda K^+$	( 6.1 ± 1.2 ) × 10 <sup>-4</sup>		781
$\Lambda K^+\pi^+\pi^-$	< 5 × 10 <sup>-4</sup>	CL=90%	637
$\Sigma^0K^+$	( 5.2 ± 0.8 ) × 10 <sup>-4</sup>		735
$\Sigma^0K^+\pi^+\pi^-$	< 2.6 × 10 <sup>-4</sup>	CL=90%	574
$\Sigma^+K^+\pi^-$	( 2.1 ± 0.6 ) × 10 <sup>-3</sup>		670
$\Sigma^+K^*(892)^0$	[r] ( 3.6 ± 1.0 ) × 10 <sup>-3</sup>		470
$\Sigma^-K^+\pi^+$	< 1.2 × 10 <sup>-3</sup>	CL=90%	664

## Doubly Cabibbo-suppressed modes

$pK^+\pi^-$	< 2.9 × 10 <sup>-4</sup>	CL=90%	823
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## Semileptonic modes

$\Lambda e^+\nu_e$	( 3.6 ± 0.4 ) %		871
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## Inclusive modes

$e^+$ anything	( 4.5 ± 1.7 ) %		—
$p e^+$ anything	( 1.8 ± 0.9 ) %		—
$p$ anything	( 50 ± 16 ) %		—
$p$ anything (no $\Lambda$ )	( 12 ± 19 ) %		—
$n$ anything	( 50 ± 16 ) %		—
$n$ anything (no $\Lambda$ )	( 29 ± 17 ) %		—
$\Lambda$ anything	( 35 ± 11 ) %	S=1.4	—
$\Sigma^\pm$ anything	[s] ( 10 ± 5 ) %		—
3prongs	( 24 ± 8 ) %		—

 $\Delta C = 1$  weak neutral current ( $C1$ ) modes, or  
Lepton Family number ( $LF$ ), or Lepton number ( $L$ ), or  
Baryon number ( $B$ ) violating modes

$p e^+ e^-$	$C1$	< 5.5 × 10 <sup>-6</sup>	CL=90%	951
$p\mu^+\mu^-$	$C1$	< 4.4 × 10 <sup>-5</sup>	CL=90%	937
$p e^+\mu^-$	$LF$	< 9.9 × 10 <sup>-6</sup>	CL=90%	947
$p e^-\mu^+$	$LF$	< 1.9 × 10 <sup>-5</sup>	CL=90%	947
$\bar{p}2e^+$	$L, B$	< 2.7 × 10 <sup>-6</sup>	CL=90%	951
$\bar{p}2\mu^+$	$L, B$	< 9.4 × 10 <sup>-6</sup>	CL=90%	937
$\bar{p}e^+\mu^+$	$L, B$	< 1.6 × 10 <sup>-5</sup>	CL=90%	947
$\Sigma^-\mu^+\mu^+$	$L$	< 7.0 × 10 <sup>-4</sup>	CL=90%	812

 $\Lambda_c(2595)^+$ 

$$I(J^P) = 0(\frac{1}{2}^-)$$

The spin-parity follows from the fact that  $\Sigma_c(2455)\pi$  decays, with little available phase space, are dominant. This assumes that  $J^P = 1/2^+$  for the  $\Sigma_c(2455)$ .

Mass  $m = 2592.25 \pm 0.28$  MeV

$m - m_{\Lambda_c^+} = 305.79 \pm 0.24$  MeV

Full width  $\Gamma = 2.6 \pm 0.6$  MeV

$\Lambda_c^+\pi\pi$  and its submode  $\Sigma_c(2455)\pi$  — the latter just barely — are the only strong decays allowed to an excited  $\Lambda_c^+$  having this mass; and the submode seems to dominate.

$\Lambda_c(2595)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda_c^+\pi^+\pi^-$	[t] —	117
$\Sigma_c(2455)^{++}\pi^-$	24 ± 7 %	†
$\Sigma_c(2455)^0\pi^+$	24 ± 7 %	†
$\Lambda_c^+\pi^+\pi^-$ 3-body	18 ± 10 %	117
$\Lambda_c^+\pi^0$	[u] not seen	258
$\Lambda_c^+\gamma$	not seen	288

 $\Lambda_c(2625)^+$ 

$$I(J^P) = 0(\frac{3}{2}^-)$$

$J^P$  has not been measured;  $\frac{3}{2}^-$  is the quark-model prediction.

Mass  $m = 2628.11 \pm 0.19$  MeV ( $S = 1.1$ )

$m - m_{\Lambda_c^+} = 341.65 \pm 0.13$  MeV ( $S = 1.1$ )

Full width  $\Gamma < 0.97$  MeV, CL = 90%

$\Lambda_c^+\pi\pi$  and its submode  $\Sigma_c(2455)\pi$  are the only strong decays allowed to an excited  $\Lambda_c^+$  having this mass.

# Baryon Summary Table

$\Lambda_c(2625)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
$\Lambda_c^+ \pi^+ \pi^-$	[t] $\approx 67\%$		184
$\Sigma_c(2455)^{++} \pi^-$	$<5$	90%	102
$\Sigma_c(2455)^0 \pi^+$	$<5$	90%	102
$\Lambda_c^+ \pi^+ \pi^-$ 3-body	large		184
$\Lambda_c^+ \pi^0$	[u] not seen		293
$\Lambda_c^+ \gamma$	not seen		319

$$\Lambda_c(2880)^+ \quad I(J^P) = 0(\frac{5}{2}^+)$$

There is some good evidence that indeed  $J^P = 5/2^+$

$$\text{Mass } m = 2881.53 \pm 0.35 \text{ MeV}$$

$$m - m_{\Lambda_c^+} = 595.1 \pm 0.4 \text{ MeV}$$

$$\text{Full width } \Gamma = 5.8 \pm 1.1 \text{ MeV}$$

$\Lambda_c(2880)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_c^+ \pi^+ \pi^-$	seen	471
$\Sigma_c(2455)^0, ++ \pi^\pm$	seen	376
$\Sigma_c(2520)^0, ++ \pi^\pm$	seen	317
$\rho D^0$	seen	316

$$\Lambda_c(2940)^+ \quad I(J^P) = 0(?^?)$$

$$\text{Mass } m = 2939.3^{+1.4}_{-1.5} \text{ MeV}$$

$$\text{Full width } \Gamma = 17^{+8}_{-6} \text{ MeV}$$

$\Lambda_c(2940)^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\rho D^0$	seen	420
$\Sigma_c(2455)^0, ++ \pi^\pm$	seen	—

$$\Sigma_c(2455) \quad I(J^P) = 1(\frac{1}{2}^+)$$

$$\Sigma_c(2455)^{++} \text{ mass } m = 2453.97 \pm 0.14 \text{ MeV}$$

$$\Sigma_c(2455)^+ \text{ mass } m = 2452.9 \pm 0.4 \text{ MeV}$$

$$\Sigma_c(2455)^0 \text{ mass } m = 2453.75 \pm 0.14 \text{ MeV}$$

$$m_{\Sigma_c^{++}} - m_{\Lambda_c^+} = 167.510 \pm 0.017 \text{ MeV}$$

$$m_{\Sigma_c^+} - m_{\Lambda_c^+} = 166.4 \pm 0.4 \text{ MeV}$$

$$m_{\Sigma_c^0} - m_{\Lambda_c^+} = 167.290 \pm 0.017 \text{ MeV}$$

$$m_{\Sigma_c^{++}} - m_{\Sigma_c^0} = 0.220 \pm 0.013 \text{ MeV}$$

$$m_{\Sigma_c^+} - m_{\Sigma_c^0} = -0.9 \pm 0.4 \text{ MeV}$$

$$\Sigma_c(2455)^{++} \text{ full width } \Gamma = 1.89^{+0.09}_{-0.18} \text{ MeV} \quad (S = 1.1)$$

$$\Sigma_c(2455)^+ \text{ full width } \Gamma < 4.6 \text{ MeV, CL} = 90\%$$

$$\Sigma_c(2455)^0 \text{ full width } \Gamma = 1.83^{+0.11}_{-0.19} \text{ MeV} \quad (S = 1.2)$$

$\Lambda_c^+ \pi$  is the only strong decay allowed to a  $\Sigma_c$  having this mass.

$\Sigma_c(2455)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_c^+ \pi$	$\approx 100\%$	94

$$\Sigma_c(2520) \quad I(J^P) = 1(\frac{3}{2}^+)$$

$J^P$  has not been measured;  $\frac{3}{2}^+$  is the quark-model prediction.

$$\Sigma_c(2520)^{++} \text{ mass } m = 2518.41^{+0.21}_{-0.19} \text{ MeV} \quad (S = 1.1)$$

$$\Sigma_c(2520)^+ \text{ mass } m = 2517.5 \pm 2.3 \text{ MeV}$$

$$\Sigma_c(2520)^0 \text{ mass } m = 2518.48 \pm 0.20 \text{ MeV} \quad (S = 1.1)$$

$$m_{\Sigma_c(2520)^{++}} - m_{\Lambda_c^+} = 231.95^{+0.17}_{-0.12} \text{ MeV} \quad (S = 1.3)$$

$$m_{\Sigma_c(2520)^+} - m_{\Lambda_c^+} = 231.0 \pm 2.3 \text{ MeV}$$

$$m_{\Sigma_c(2520)^0} - m_{\Lambda_c^+} = 232.02^{+0.15}_{-0.14} \text{ MeV} \quad (S = 1.3)$$

$$m_{\Sigma_c(2520)^{++}} - m_{\Sigma_c(2520)^0} = 0.01 \pm 0.15 \text{ MeV}$$

$$\Sigma_c(2520)^{++} \text{ full width } \Gamma = 14.78^{+0.30}_{-0.40} \text{ MeV}$$

$$\Sigma_c(2520)^+ \text{ full width } \Gamma < 17 \text{ MeV, CL} = 90\%$$

$$\Sigma_c(2520)^0 \text{ full width } \Gamma = 15.3^{+0.4}_{-0.5} \text{ MeV}$$

$\Lambda_c^+ \pi$  is the only strong decay allowed to a  $\Sigma_c$  having this mass.

$\Sigma_c(2520)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_c^+ \pi$	$\approx 100\%$	179

$$\Sigma_c(2800) \quad I(J^P) = 1(?^?)$$

$$\Sigma_c(2800)^{++} \text{ mass } m = 2801^{+4}_{-6} \text{ MeV}$$

$$\Sigma_c(2800)^+ \text{ mass } m = 2792^{+14}_{-5} \text{ MeV}$$

$$\Sigma_c(2800)^0 \text{ mass } m = 2806^{+5}_{-7} \text{ MeV} \quad (S = 1.3)$$

$$m_{\Sigma_c(2800)^{++}} - m_{\Lambda_c^+} = 514^{+4}_{-6} \text{ MeV}$$

$$m_{\Sigma_c(2800)^+} - m_{\Lambda_c^+} = 505^{+14}_{-5} \text{ MeV}$$

$$m_{\Sigma_c(2800)^0} - m_{\Lambda_c^+} = 519^{+5}_{-7} \text{ MeV} \quad (S = 1.3)$$

$$\Sigma_c(2800)^{++} \text{ full width } \Gamma = 75^{+22}_{-17} \text{ MeV}$$

$$\Sigma_c(2800)^+ \text{ full width } \Gamma = 62^{+60}_{-40} \text{ MeV}$$

$$\Sigma_c(2800)^0 \text{ full width } \Gamma = 72^{+22}_{-15} \text{ MeV}$$

$\Sigma_c(2800)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_c^+ \pi$	seen	443

$$\Xi_c^+ \quad I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

$J^P$  has not been measured;  $\frac{1}{2}^+$  is the quark-model prediction.

$$\text{Mass } m = 2467.93^{+0.28}_{-0.40} \text{ MeV}$$

$$\text{Mean life } \tau = (442 \pm 26) \times 10^{-15} \text{ s} \quad (S = 1.3)$$

$$c\tau = 132 \mu\text{m}$$

$\Xi_c^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	$\rho$ (MeV/c)
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**No absolute branching fractions have been measured.**  
The following are branching *ratios* relative to  $\Xi^- 2\pi^+$ .

**Cabibbo-favored ( $S = -2$ ) decays — relative to  $\Xi^- 2\pi^+$**

$\rho 2K_S^0$		0.087 $\pm$ 0.021	767
$\Lambda \bar{K}^0 \pi^+$		—	852
$\Sigma(1385)^+ \bar{K}^0$	[r]	1.0 $\pm$ 0.5	746
$\Lambda K^- 2\pi^+$		0.323 $\pm$ 0.033	787
$\Lambda \bar{K}^*(892)^0 \pi^+$	[r]	$<0.16$	90% 608
$\Sigma(1385)^+ K^- \pi^+$	[r]	$<0.23$	90% 678
$\Sigma^+ K^- \pi^+$		0.94 $\pm$ 0.10	811
$\Sigma^+ \bar{K}^*(892)^0$	[r]	0.81 $\pm$ 0.15	658
$\Sigma^0 K^- 2\pi^+$		0.27 $\pm$ 0.12	735
$\Xi^0 \pi^+$		0.55 $\pm$ 0.16	877
$\Xi^- 2\pi^+$		<b>DEFINED AS 1</b>	851
$\Xi(1530)^0 \pi^+$	[r]	$<0.10$	90% 750
$\Xi^0 \pi^+ \pi^0$		2.3 $\pm$ 0.7	856
$\Xi^0 \pi^- 2\pi^+$		1.7 $\pm$ 0.5	818
$\Xi^0 e^+ \nu_e$		2.3 $\pm$ 0.7 — 0.8	884
$\Omega^- K^+ \pi^+$		0.07 $\pm$ 0.04	399

**Cabibbo-suppressed decays — relative to  $\Xi^- 2\pi^+$**

$\rho K^- \pi^+$		0.21 $\pm$ 0.04	944
$\rho \bar{K}^*(892)^0$	[r]	0.116 $\pm$ 0.030	828
$\Sigma^+ \pi^+ \pi^-$		0.48 $\pm$ 0.20	922
$\Sigma^- 2\pi^+$		0.18 $\pm$ 0.09	918
$\Sigma^+ K^+ K^-$		0.15 $\pm$ 0.06	580
$\Sigma^+ \phi$	[r]	$<0.11$	90% 549
$\Xi(1690)^0 K^+, \Xi^0 \rightarrow$		$<0.05$	90% 501
$\Sigma^+ K^-$			

$$\Xi_c^0 \quad I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

$J^P$  has not been measured;  $\frac{1}{2}^+$  is the quark-model prediction.

$$\text{Mass } m = 2470.85^{+0.28}_{-0.40} \text{ MeV}$$

$$m_{\Xi_c^0} - m_{\Xi_c^+} = 2.93 \pm 0.24 \text{ MeV}$$

$$\text{Mean life } \tau = (112^{+13}_{-10}) \times 10^{-15} \text{ s}$$

$$c\tau = 33.6 \mu\text{m}$$

## Baryon Summary Table

## Decay asymmetry parameters

$$\Xi^- \pi^+ \quad \alpha = -0.6 \pm 0.4$$

No absolute branching fractions have been measured. Several measurements of ratios of fractions may be found in the Listings that follow.

$\Xi_c^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
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No absolute branching fractions have been measured.  
The following are branching *ratios* relative to  $\Xi^- \pi^+$ .

Cabibbo-favored (S = -2) decays — relative to  $\Xi^- \pi^+$ 

$p K^- K^- \pi^+$	0.34 $\pm$ 0.04	676
$p K^- \bar{K}^*(892)^0$	0.21 $\pm$ 0.05	413
$p K^- K^- \pi^+$ (no $\bar{K}^{*0}$ )	0.21 $\pm$ 0.04	676
$\Lambda K_S^0$	0.210 $\pm$ 0.028	906
$\Lambda K^- \pi^+$	1.07 $\pm$ 0.14	856
$\Lambda \bar{K}^0 \pi^+ \pi^-$	seen	787
$\Lambda K^- \pi^+ \pi^+ \pi^-$	seen	703
$\Xi^- \pi^+$	DEFINED AS 1	875
$\Xi^- \pi^+ \pi^+ \pi^-$	3.3 $\pm$ 1.4	816
$\Omega^- K^+$	0.297 $\pm$ 0.024	522
$\Xi^- e^+ \nu_e$	3.1 $\pm$ 1.1	882
$\Xi^- \ell^+$ anything	1.0 $\pm$ 0.5	—

Cabibbo-suppressed decays — relative to  $\Xi^- \pi^+$ 

$\Xi^- K^+$	0.028 $\pm$ 0.006	790
$\Lambda K^+ K^-$ (no $\phi$ )	0.029 $\pm$ 0.007	648
$\Lambda \phi$	0.034 $\pm$ 0.007	621

$\Xi_c^{'+}$	$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$
$J^P$ has not been measured; $\frac{1}{2}^+$ is the quark-model prediction.	
Mass $m = 2575.7 \pm 3.0$ MeV	
$m_{\Xi_c^{'+}} - m_{\Xi_c^+} = 107.8 \pm 3.0$ MeV	
The $\Xi_c^{'+} - \Xi_c^+$ mass difference is too small for any strong decay to occur.	

$\Xi_c^{'+}$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_c^{'+} \gamma$	seen	106

$\Xi_c^0$	$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$
$J^P$ has not been measured; $\frac{1}{2}^+$ is the quark-model prediction.	
Mass $m = 2577.9 \pm 2.9$ MeV	
$m_{\Xi_c^0} - m_{\Xi_c^+} = 107.0 \pm 2.9$ MeV	
The $\Xi_c^0 - \Xi_c^+$ mass difference is too small for any strong decay to occur.	

$\Xi_c^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_c^0 \gamma$	seen	105

$\Xi_c(2645)$	$I(J^P) = \frac{1}{2}(\frac{3}{2}^+)$
$J^P$ has not been measured; $\frac{3}{2}^+$ is the quark-model prediction.	
$\Xi_c(2645)^+$ mass $m = 2645.9 \pm 0.5$ MeV (S = 1.1)	
$\Xi_c(2645)^0$ mass $m = 2645.9 \pm 0.5$ MeV	
$m_{\Xi_c(2645)^+} - m_{\Xi_c^0} = 175.0 \pm 0.6$ MeV (S = 1.1)	
$m_{\Xi_c(2645)^0} - m_{\Xi_c^+} = 178.0 \pm 0.6$ MeV	
$m_{\Xi_c(2645)^+} - m_{\Xi_c(2645)^0} = 0.0 \pm 0.5$ MeV	
$\Xi_c(2645)^+$ full width $\Gamma = 2.6 \pm 0.4$ MeV	
$\Xi_c(2645)^0$ full width $\Gamma < 5.5$ MeV, CL = 90%	
$\Xi_c \pi$ is the only strong decay allowed to a $\Xi_c$ resonance having this mass.	

$\Xi_c(2645)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_c^0 \pi^+$	seen	102
$\Xi_c^+ \pi^-$	seen	107

$\Xi_c(2790)$	$I(J^P) = \frac{1}{2}(\frac{1}{2}^-)$
$J^P$ has not been measured; $\frac{1}{2}^-$ is the quark-model prediction.	
$\Xi_c(2790)^+$ mass = 2789.1 $\pm$ 3.2 MeV	
$\Xi_c(2790)^0$ mass = 2791.9 $\pm$ 3.3 MeV	
$m_{\Xi_c(2790)^+} - m_{\Xi_c^0} = 318.2 \pm 3.2$ MeV	
$m_{\Xi_c(2790)^0} - m_{\Xi_c^+} = 324.0 \pm 3.3$ MeV	
$\Xi_c(2790)^+$ width < 15 MeV, CL = 90%	
$\Xi_c(2790)^0$ width < 12 MeV, CL = 90%	

$\Xi_c(2790)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_c' \pi$	seen	159

$\Xi_c(2815)$	$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$
$J^P$ has not been measured; $\frac{3}{2}^-$ is the quark-model prediction.	
$\Xi_c(2815)^+$ mass $m = 2816.6 \pm 0.9$ MeV	
$\Xi_c(2815)^0$ mass $m = 2819.6 \pm 1.2$ MeV	
$m_{\Xi_c(2815)^+} - m_{\Xi_c^+} = 348.7 \pm 0.9$ MeV	
$m_{\Xi_c(2815)^0} - m_{\Xi_c^0} = 348.8 \pm 1.2$ MeV	
$m_{\Xi_c(2815)^+} - m_{\Xi_c(2815)^0} = -3.0 \pm 1.3$ MeV	
$\Xi_c(2815)^+$ full width $\Gamma < 3.5$ MeV, CL = 90%	
$\Xi_c(2815)^0$ full width $\Gamma < 6.5$ MeV, CL = 90%	
The $\Xi_c \pi \pi$ modes are consistent with being entirely via $\Xi_c(2645) \pi$ .	

$\Xi_c(2815)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_c^+ \pi^+ \pi^-$	seen	196
$\Xi_c^0 \pi^+ \pi^-$	seen	191

$\Xi_c(2970)$ was $\Xi_c(2980)$	$I(J^P) = \frac{1}{2}(?)^?$
$\Xi_c(2970)^+$ $m = 2970.7 \pm 2.2$ MeV (S = 1.5)	
$\Xi_c(2970)^0$ $m = 2968.0 \pm 2.6$ MeV (S = 1.2)	
$\Xi_c(2970)^+$ width $\Gamma = 17.9 \pm 3.5$ MeV	
$\Xi_c(2970)^0$ width $\Gamma = 20 \pm 7$ MeV (S = 1.3)	

$\Xi_c(2970)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_c^+ \bar{K} \pi$	seen	231
$\Sigma_c(2455) \bar{K}$	seen	134
$\Lambda_c^+ \bar{K}$	not seen	414
$\Xi_c 2\pi$	seen	385
$\Xi_c(2645) \pi$	seen	277

$\Xi_c(3055)$	$I(J^P) = ?(?)^?$
Mass $m = 3055.1 \pm 1.7$ MeV (S = 1.5)	
Full width $\Gamma = 11 \pm 4$ MeV	

$\Xi_c(3080)$	$I(J^P) = \frac{1}{2}(?)^?$
$\Xi_c(3080)^+$ $m = 3076.94 \pm 0.28$ MeV	
$\Xi_c(3080)^0$ $m = 3079.9 \pm 1.4$ MeV (S = 1.3)	
$\Xi_c(3080)^+$ width $\Gamma = 4.3 \pm 1.5$ MeV (S = 1.3)	
$\Xi_c(3080)^0$ width $\Gamma = 5.6 \pm 2.2$ MeV	

# Baryon Summary Table

$\Xi_c(3080)$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda_c^+ \bar{K} \pi$	seen	415
$\Sigma_c(2455) \bar{K}$	seen	342
$\Sigma_c(2455) \bar{K} + \Sigma_c(2520) \bar{K}$	seen	–
$\Lambda_c^+ \bar{K}$	not seen	536
$\Lambda_c^+ \bar{K} \pi^+ \pi^-$	not seen	143

$\Omega_c^0$

$I(J^P) = 0(\frac{1}{2}^+)$

$J^P$  has not been measured;  $\frac{1}{2}^+$  is the quark-model prediction.

Mass  $m = 2695.2 \pm 1.7$  MeV ( $S = 1.3$ )

Mean life  $\tau = (69 \pm 12) \times 10^{-15}$  s

$c\tau = 21 \mu\text{m}$

No absolute branching fractions have been measured.

$\Omega_c^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Sigma^+ K^- K^- \pi^+$	seen	689
$\Xi^0 K^- \pi^+$	seen	901
$\Xi^- K^- \pi^+ \pi^+$	seen	830
$\Omega^- e^+ \nu_e$	seen	829
$\Omega^- \pi^+$	seen	821
$\Omega^- \pi^+ \pi^0$	seen	797
$\Omega^- \pi^- \pi^+ \pi^+$	seen	753

$\Omega_c(2770)^0$

$I(J^P) = 0(\frac{3}{2}^+)$

$J^P$  has not been measured;  $\frac{3}{2}^+$  is the quark-model prediction.

Mass  $m = 2765.9 \pm 2.0$  MeV ( $S = 1.2$ )

$m_{\Omega_c(2770)^0} - m_{\Omega_c^0} = 70.7^{+0.8}_{-0.9}$  MeV

The  $\Omega_c(2770)^0 - \Omega_c^0$  mass difference is too small for any strong decay to occur.

$\Omega_c(2770)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Omega_c^0 \gamma$	presumably 100%	70

BOTTOM BARYONS

$(B = -1)$

$\Lambda_b^0 = udb, \Xi_b^0 = usb, \Xi_b^- = dsb, \Omega_b^- = ssb$

$\Lambda_b^0$

$I(J^P) = 0(\frac{1}{2}^+)$

$I(J^P)$  not yet measured;  $0(\frac{1}{2}^+)$  is the quark model prediction.

Mass  $m = 5619.51 \pm 0.23$  MeV

$m_{\Lambda_b^0} - m_{B^0} = 339.2 \pm 1.4$  MeV

$m_{\Lambda_b^0} - m_{B^+} = 339.72 \pm 0.28$  MeV

Mean life  $\tau = (1.466 \pm 0.010) \times 10^{-12}$  s

$c\tau = 439.5 \mu\text{m}$

$A_{CP}(\Lambda_b \rightarrow p\pi^-) = 0.06 \pm 0.07$

$A_{CP}(\Lambda_b \rightarrow pK^-) = 0.00 \pm 0.19$  ( $S = 2.4$ )

$A_{CP}(\Lambda_b \rightarrow p\bar{K}^0\pi^-) = 0.22 \pm 0.13$

$\Delta A_{CP}(J/\psi p\pi^-/K^-) \equiv A_{CP}(J/\psi p\pi^-) - A_{CP}(J/\psi pK^-)$   
 $= (5.7 \pm 2.7) \times 10^{-2}$

$\alpha$  decay parameter for  $\Lambda_b \rightarrow J/\psi \Lambda = 0.18 \pm 0.13$

$A_{FB}^{\ell}(\mu\mu)$  in  $\Lambda_b \rightarrow \Lambda\mu^+\mu^- = -0.05 \pm 0.09$

$A_{FB}^h(p\pi)$  in  $\Lambda_b \rightarrow \Lambda(p\pi)\mu^+\mu^- = -0.29 \pm 0.08$

$f_L(\mu\mu)$  longitudinal polarization fraction in  $\Lambda_b \rightarrow \Lambda\mu^+\mu^- =$   
 $0.61^{+0.11}_{-0.14}$

The branching fractions  $B(b\text{-baryon} \rightarrow \Lambda \ell^- \bar{\nu}_\ell \text{anything})$  and  $B(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})$  are not pure measurements because the underlying measured products of these with  $B(b \rightarrow b\text{-baryon})$  were used to determine  $B(b \rightarrow b\text{-baryon})$ , as described in the note “Production and Decay of  $b$ -Flavored Hadrons.”

For inclusive branching fractions, e.g.,  $\Lambda_b \rightarrow \bar{\Lambda}_c \text{anything}$ , the values usually are multiplicities, not branching fractions. They can be greater than one.

$\Lambda_b^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor / Confidence level	$p$ (MeV/c)
$J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0)$	$(5.8 \pm 0.8) \times 10^{-5}$		1740
$p D^0 \pi^-$	$(6.4 \pm 0.7) \times 10^{-4}$		2370
$p D^0 K^-$	$(4.7 \pm 0.8) \times 10^{-5}$		2269
$p J/\psi \pi^-$	$(2.6^{+0.5}_{-0.4}) \times 10^{-5}$		1755
$p J/\psi K^-$	$(3.2^{+0.6}_{-0.5}) \times 10^{-4}$		1589
$P_c(4380)^+ K^-, P_c \rightarrow p J/\psi$	[v] $(2.7 \pm 1.4) \times 10^{-5}$		–
$P_c(4450)^+ K^-, P_c \rightarrow p J/\psi$	[v] $(1.3 \pm 0.4) \times 10^{-5}$		–
$p \bar{K}^0 \pi^-$	$(1.3 \pm 0.4) \times 10^{-5}$		2693
$p K^0 K^-$	$< 3.5 \times 10^{-6}$	CL=90%	2639
$\Lambda_c^+ \pi^-$	$(4.9 \pm 0.4) \times 10^{-3}$	S=1.2	2342
$\Lambda_c^+ K^-$	$(3.59 \pm 0.30) \times 10^{-4}$	S=1.2	2314
$\Lambda_c^+ a_1(1260)^-$	seen		2153
$\Lambda_c^+ D^-$	$(4.6 \pm 0.6) \times 10^{-4}$		1886
$\Lambda_c^+ D_s^-$	$(1.10 \pm 0.10) \%$		1833
$\Lambda_c^+ \pi^+ \pi^- \pi^-$	$(7.7 \pm 1.1) \times 10^{-3}$	S=1.1	2323
$\Lambda_c(2595)^+ \pi^-, \Lambda_c(2595)^+ \rightarrow \Lambda_c^+ \pi^+ \pi^-$	$(3.4 \pm 1.5) \times 10^{-4}$		2210
$\Lambda_c(2625)^+ \pi^-, \Lambda_c(2625)^+ \rightarrow \Lambda_c^+ \pi^+ \pi^-$	$(3.3 \pm 1.3) \times 10^{-4}$		2193
$\Sigma_c(2455)^0 \pi^+ \pi^-, \Sigma_c^0 \rightarrow \Lambda_c^+ \pi^-$	$(5.7 \pm 2.2) \times 10^{-4}$		2265
$\Sigma_c(2455)^{++} \pi^- \pi^-, \Sigma_c^{++} \rightarrow \Lambda_c^+ \pi^+$	$(3.2 \pm 1.6) \times 10^{-4}$		2265
$\Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything}$	[x] $(10.3 \pm 2.2) \%$		–
$\Lambda_c^+ \ell^- \bar{\nu}_\ell$	$(6.2^{+1.4}_{-1.3}) \%$		2345
$\Lambda_c^+ \pi^+ \pi^- \ell^- \bar{\nu}_\ell$	$(5.6 \pm 3.1) \%$		2335
$\Lambda_c(2595)^+ \ell^- \bar{\nu}_\ell$	$(7.9^{+4.0}_{-3.5}) \times 10^{-3}$		2212
$\Lambda_c(2625)^+ \ell^- \bar{\nu}_\ell$	$(1.3^{+0.6}_{-0.5}) \%$		2195
$p h^-$	[y] $< 2.3 \times 10^{-5}$	CL=90%	2730
$p \pi^-$	$(4.2 \pm 0.8) \times 10^{-6}$		2730
$p K^-$	$(5.1 \pm 1.0) \times 10^{-6}$		2708
$p D_s^-$	$< 4.8 \times 10^{-4}$	CL=90%	2364
$p \mu^- \bar{\nu}_\mu$	$(4.1 \pm 1.0) \times 10^{-4}$		2730
$\Lambda \mu^+ \mu^-$	$(1.08 \pm 0.28) \times 10^{-6}$		2695
$\Lambda \gamma$	$< 1.3 \times 10^{-3}$	CL=90%	2699
$\Lambda^0 \eta$	$(9^{+7}_{-5}) \times 10^{-6}$		–
$\Lambda^0 \eta'(958)$	$< 3.1 \times 10^{-6}$	CL=90%	–

$\Lambda_b(5912)^0$

$J^P = \frac{1}{2}^-$

Mass  $m = 5912.11 \pm 0.26$  MeV

Full width  $\Gamma < 0.66$  MeV, CL = 90%

$\Lambda_b(5912)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda_b^0 \pi^+ \pi^-$	seen	86

$\Lambda_b(5920)^0$

$J^P = \frac{3}{2}^-$

Mass  $m = 5919.81 \pm 0.23$  MeV

Full width  $\Gamma < 0.63$  MeV, CL = 90%

$\Lambda_b(5920)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$p$ (MeV/c)
$\Lambda_b^0 \pi^+ \pi^-$	seen	108

## Baryon Summary Table

 **$\Sigma_b$** 

$$I(J^P) = 1(\frac{1}{2}^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m(\Sigma_b^+) = 5811.3 \pm 1.9 \text{ MeV}$$

$$\text{Mass } m(\Sigma_b^-) = 5815.5 \pm 1.8 \text{ MeV}$$

$$m_{\Sigma_b^+} - m_{\Sigma_b^-} = -4.2 \pm 1.1 \text{ MeV}$$

$$\Gamma(\Sigma_b^+) = 9.7^{+4.0}_{-3.0} \text{ MeV}$$

$$\Gamma(\Sigma_b^-) = 4.9^{+3.3}_{-2.4} \text{ MeV}$$

$\Sigma_b$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_b^0 \pi$	dominant	134

 **$\Sigma_b^*$** 

$$I(J^P) = 1(\frac{3}{2}^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m(\Sigma_b^{*+}) = 5832.1 \pm 1.9 \text{ MeV}$$

$$\text{Mass } m(\Sigma_b^{*-}) = 5835.1 \pm 1.9 \text{ MeV}$$

$$m_{\Sigma_b^{*+}} - m_{\Sigma_b^{*-}} = -3.0^{+1.0}_{-0.9} \text{ MeV}$$

$$\Gamma(\Sigma_b^{*+}) = 11.5 \pm 2.8 \text{ MeV}$$

$$\Gamma(\Sigma_b^{*-}) = 7.5 \pm 2.3 \text{ MeV}$$

$$m_{\Sigma_b^*} - m_{\Sigma_b} = 21.2 \pm 2.0 \text{ MeV}$$

$\Sigma_b^*$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Lambda_b^0 \pi$	dominant	161

 **$\Xi_b^0, \Xi_b^-$** 

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

$I, J, P$  need confirmation.

$$m(\Xi_b^-) = 5794.5 \pm 1.4 \text{ MeV} \quad (S = 4.0)$$

$$m(\Xi_b^0) = 5791.9 \pm 0.5 \text{ MeV}$$

$$m_{\Xi_b^-} - m_{\Lambda_b^0} = 177.9 \pm 0.9 \text{ MeV} \quad (S = 2.1)$$

$$m_{\Xi_b^0} - m_{\Lambda_b^0} = 172.5 \pm 0.4 \text{ MeV}$$

$$m_{\Xi_b^-} - m_{\Xi_b^0} = 5.9 \pm 0.6 \text{ MeV}$$

$$\text{Mean life } \tau_{\Xi_b^-} = (1.560 \pm 0.040) \times 10^{-12} \text{ s}$$

$$\text{Mean life } \tau_{\Xi_b^0} = (1.464 \pm 0.031) \times 10^{-12} \text{ s}$$

$\Xi_b$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	$\rho$ (MeV/c)
$\Xi_b^- \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X \times B(\bar{b} \rightarrow \Xi_b^-)$	$(3.9 \pm 1.2) \times 10^{-4}$	S=1.4	—
$\Xi_b^- \rightarrow J/\psi \Xi^- \times B(b \rightarrow \Xi_b^-)$	$(1.02^{+0.26}_{-0.21}) \times 10^{-5}$		1782
$\Xi_b^0 \rightarrow p D^0 K^- \times B(\bar{b} \rightarrow \Xi_b^0)$	$(1.7 \pm 0.6) \times 10^{-6}$		2374
$\Xi_b^0 \rightarrow p \bar{K}^0 \pi^- \times B(\bar{b} \rightarrow \Xi_b^0)$	$< 1.6 \times 10^{-6}$	CL=90%	2783
$\Xi_b^0/B(\bar{b} \rightarrow B^0)$			
$\Xi_b^0 \rightarrow p K^0 K^- \times B(\bar{b} \rightarrow \Xi_b^0)$	$< 1.1 \times 10^{-6}$	CL=90%	2730
$\Xi_b^0/B(\bar{b} \rightarrow B^0)$			
$\Xi_b^0 \rightarrow \Lambda_c^+ K^- \times B(\bar{b} \rightarrow \Xi_b^0)$	$(6 \pm 4) \times 10^{-7}$		2416
$\Xi_b^- \rightarrow \Lambda_b^0 \pi^- \times B(b \rightarrow \Xi_b^-)$	$(5.7 \pm 2.0) \times 10^{-4}$		100
$\Xi_b^-/B(b \rightarrow \Lambda_b^0)$			

 **$\Xi'_b(5935)^-$** 

$$J^P = \frac{1}{2}^+$$

$$\text{Mass } m = 5935.02 \pm 0.05 \text{ MeV}$$

$$m_{\Xi'_b(5935)^-} - m_{\Xi_b^0} - m_{\pi^-} = 3.653 \pm 0.019 \text{ MeV}$$

$$\text{Full width } \Gamma < 0.08 \text{ MeV, CL} = 95\%$$

$\Xi'_b(5935)^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_b^0 \pi^- \times B(\bar{b} \rightarrow \Xi'_b(5935)^-)/B(\bar{b} \rightarrow \Xi_b^0)$	$(11.8 \pm 1.8) \%$	31

 **$\Xi_b(5945)^0$** 

$$J^P = \frac{3}{2}^+$$

$$\text{Mass } m = 5948.9 \pm 1.6 \text{ MeV}$$

$$\text{Full width } \Gamma = 2.1 \pm 1.7 \text{ MeV}$$

$\Xi_b(5945)^0$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_b^- \pi^+$	seen	71

 **$\Xi_b^*(5955)^-$** 

$$J^P = \frac{3}{2}^+$$

$$\text{Mass } m = 5955.33 \pm 0.13 \text{ MeV}$$

$$m_{\Xi_b^*(5955)^-} - m_{\Xi_b^0} - m_{\pi^-} = 23.96 \pm 0.13 \text{ MeV}$$

$$\text{Full width } \Gamma = 1.65 \pm 0.33 \text{ MeV}$$

$\Xi_b^*(5955)^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$\Xi_b^0 \pi^- \times B(\bar{b} \rightarrow \Xi_b^*(5955)^-)/B(\bar{b} \rightarrow \Xi_b^0)$	$(20.7 \pm 3.5) \%$	84

 **$\Omega_b^-$** 

$$I(J^P) = 0(\frac{1}{2}^+)$$

$I, J, P$  need confirmation.

$$\text{Mass } m = 6046.4 \pm 1.9 \text{ MeV}$$

$$m_{\Omega_b^-} - m_{\Lambda_b^0} = 426.4 \pm 2.2 \text{ MeV}$$

$$\text{Mean life } \tau = (1.57^{+0.23}_{-0.20}) \times 10^{-12} \text{ s}$$

$\Omega_b^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$J/\psi \Omega^- \times B(b \rightarrow \Omega_b^-)$	$(2.9^{+1.1}_{-0.8}) \times 10^{-6}$	1806

 **$b$ -baryon ADMIXTURE ( $\Lambda_b, \Xi_b, \Sigma_b, \Omega_b$ )**Mean life  $\tau$ 

These branching fractions are actually an average over weakly decaying  $b$ -baryons weighted by their production rates at the LHC, LEP, and Tevatron, branching ratios, and detection efficiencies. They scale with the  $b$ -baryon production fraction  $B(b \rightarrow b\text{-baryon})$ .

The branching fractions  $B(b\text{-baryon} \rightarrow \Lambda \ell^- \bar{\nu}_\ell \text{anything})$  and  $B(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})$  are not pure measurements because the underlying measured products of these with  $B(b \rightarrow b\text{-baryon})$  were used to determine  $B(b \rightarrow b\text{-baryon})$ , as described in the note "Production and Decay of  $b$ -Flavored Hadrons."

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm \text{anything}$ , the values usually are multiplicities, not branching fractions. They can be greater than one.

$b$ -baryon ADMIXTURE DECAY MODES ( $\Lambda_b, \Xi_b, \Sigma_b, \Omega_b$ )	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
$p \mu^- \bar{\nu}$ anything	$(5.5^{+2.2}_{-1.9}) \%$	—
$p \ell \bar{\nu}_\ell$ anything	$(5.3 \pm 1.2) \%$	—
$p$ anything	$(66 \pm 21) \%$	—
$\Lambda \ell^- \bar{\nu}_\ell$ anything	$(3.6 \pm 0.6) \%$	—
$\Lambda \ell^+ \nu_\ell$ anything	$(3.0 \pm 0.8) \%$	—
$\Lambda$ anything	$(37 \pm 7) \%$	—
$\Xi^- \ell^- \bar{\nu}_\ell$ anything	$(6.2 \pm 1.6) \times 10^{-3}$	—

**EXOTIC BARYONS** **$P_c(4380)^+$** 

$$\text{Mass } m = 4380 \pm 30 \text{ MeV}$$

$$\text{Full width } \Gamma = 205 \pm 90 \text{ MeV}$$

Mode	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
Mode		
$J/\psi p$	seen	741

 **$P_c(4450)^+$** 

$$\text{Mass } m = 4449.8 \pm 3.0 \text{ MeV}$$

$$\text{Full width } \Gamma = 39 \pm 20 \text{ MeV}$$

Mode	Fraction ( $\Gamma_i/\Gamma$ )	$\rho$ (MeV/c)
Mode		
$J/\psi p$	seen	820

# Baryon Summary Table

## NOTES

This Summary Table only includes established baryons. The Particle Listings include evidence for other baryons. The masses, widths, and branching fractions for the resonances in this Table are Breit-Wigner parameters, but pole positions are also given for most of the  $N$  and  $\Delta$  resonances.

For most of the resonances, the parameters come from various partial-wave analyses of more or less the same sets of data, and it is not appropriate to treat the results of the analyses as independent or to average them together. Furthermore, the systematic errors on the results are not well understood. Thus, we usually only give ranges for the parameters. We then also give a best guess for the mass (as part of the name of the resonance) and for the width. The *Note on  $N$  and  $\Delta$  Resonances* and the *Note on  $\Lambda$  and  $\Sigma$  Resonances* in the Particle Listings review the partial-wave analyses.

When a quantity has “(S = ...)” to its right, the error on the quantity has been enlarged by the “scale factor” S, defined as  $S = \sqrt{\chi^2/(N-1)}$ , where  $N$  is the number of measurements used in calculating the quantity. We do this when  $S > 1$ , which often indicates that the measurements are inconsistent. When  $S > 1.25$ , we also show in the Particle Listings an ideogram of the measurements. For more about S, see the Introduction.

A decay momentum  $p$  is given for each decay mode. For a 2-body decay,  $p$  is the momentum of each decay product in the rest frame of the decaying particle. For a 3-or-more-body decay,  $p$  is the largest momentum any of the products can have in this frame. For any resonance, the *nominal* mass is used in calculating  $p$ . A dagger (“†”) in this column indicates that the mode is forbidden when the nominal masses of resonances are used, but is in fact allowed due to the nonzero widths of the resonances.

- [a] The masses of the  $p$  and  $n$  are most precisely known in u (unified atomic mass units). The conversion factor to MeV,  $1\text{ u} = 931.494061(21)\text{ MeV}$ , is less well known than are the masses in u.
- [b] The  $|m_p - m_{\bar{p}}|/m_p$  and  $|q_p + q_{\bar{p}}|/e$  are not independent, and both use the more precise measurement of  $|q_{\bar{p}}/m_{\bar{p}}|/(q_p/m_p)$ .
- [c] The limit is from neutrality-of-matter experiments; it assumes  $q_n = q_p + q_e$ . See also the charge of the neutron.
- [d] The  $\mu p$  and  $e p$  values for the charge radius are much too different to average them. The disagreement is not yet understood.
- [e] There is a lot of disagreement about the value of the proton magnetic charge radius. See the Listings.
- [f] The first limit is for  $p \rightarrow$  anything or “disappearance” modes of a bound proton. The second entry, a rough range of limits, assumes the dominant decay modes are among those investigated. For antiprotons the best limit, inferred from the observation of cosmic ray  $\bar{p}$ ’s is  $\tau_{\bar{p}} > 10^7$  yr, the cosmic-ray storage time, but this limit depends on a number of assumptions. The best direct observation of stored antiprotons gives  $\tau_{\bar{p}}/B(\bar{p} \rightarrow e^- \gamma) > 7 \times 10^5$  yr.

[g] There is some controversy about whether nuclear physics and model dependence complicate the analysis for bound neutrons (from which the best limit comes). The first limit here is from reactor experiments with free neutrons.

[h] Lee and Yang in 1956 proposed the existence of a mirror world in an attempt to restore global parity symmetry—thus a search for oscillations between the two worlds. Oscillations between the worlds would be maximal when the magnetic fields  $B$  and  $B'$  were equal. The limit for any  $B'$  in the range 0 to  $12.5\text{ }\mu\text{T}$  is  $>12\text{ s}$  (95% CL).

[i] The parameters  $g_A, g_V$ , and  $g_{WM}$  for semileptonic modes are defined by  $\bar{B}_f[\gamma_\lambda(g_V + g_A\gamma_5) + i(g_{WM}/m_{B_i})\sigma_{\lambda\nu}q^\nu]B_i$ , and  $\phi_{AV}$  is defined by  $g_A/g_V = |g_A/g_V|e^{i\phi_{AV}}$ . See the “Note on Baryon Decay Parameters” in the neutron Particle Listings.

[j] Time-reversal invariance requires this to be  $0^\circ$  or  $180^\circ$ .

[k] This coefficient is zero if time invariance is not violated.

[l] This limit is for  $\gamma$  energies between 15 and 340 keV.

[n] The decay parameters  $\gamma$  and  $\Delta$  are calculated from  $\alpha$  and  $\phi$  using

$$\gamma = \sqrt{1-\alpha^2} \cos\phi, \quad \tan\Delta = -\frac{1}{\alpha} \sqrt{1-\alpha^2} \sin\phi.$$

See the “Note on Baryon Decay Parameters” in the neutron Particle Listings.

[o] See the Listings for the pion momentum range used in this measurement.

[p] The error given here is only an educated guess. It is larger than the error on the weighted average of the published values.

[q] A theoretical value using QED.

[r] This branching fraction includes all the decay modes of the final-state resonance.

[s] The value is for the sum of the charge states or particle/antiparticle states indicated.

[t] See AALTONEN 11H, Fig. 8, for the calculated ratio of  $\Lambda_c^+ \pi^0 \pi^0$  and  $\Lambda_c^+ \pi^+ \pi^-$  partial widths as a function of the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass difference. At our value of the mass difference, the ratio is about 4.

[u] A test that the isospin is indeed 0, so that the particle is indeed a  $\Lambda_c^+$ .

[v]  $P_c^+$  is a pentaquark-charmonium state.

[x] Not a pure measurement. See note at head of  $\Lambda_b^0$  Decay Modes.

[y] Here  $h^-$  means  $\pi^-$  or  $K^-$ .

## Searches Summary Table

# SEARCHES FOR MONOPOLES, SUPERSYMMETRY, TECHNICOLOR, COMPOSITENESS, EXTRA DIMENSIONS, etc.

## Magnetic Monopole Searches

Isolated supermassive monopole candidate events have not been confirmed. The most sensitive experiments obtain negative results.

Best cosmic-ray supermassive monopole flux limit:

$$< 1.4 \times 10^{-16} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1} \quad \text{for } 1.1 \times 10^{-4} < \beta < 1$$

## Supersymmetric Particle Searches

Presently all supersymmetric mass bounds are model dependent. This table contains a selection of bounds indicating the range of possibilities. For a more extensive set of cases consult the detailed listings.

The limits are based on the Minimal Supersymmetric Standard Model (MSSM) with additional assumptions as follows:

1)  $\tilde{\chi}_1^0$  is lightest supersymmetric particle; 2)  $R$ -parity is conserved;

See the Particle Listings for a Note giving details of supersymmetry.

$\tilde{\chi}_i^0$  — neutralinos (mixtures of  $\tilde{\gamma}$ ,  $\tilde{Z}^0$ , and  $\tilde{H}_i^0$ )

Mass  $m_{\tilde{\chi}_1^0} > 0$  GeV, CL = 95%

[general MSSM, non-universal gaugino masses]

Mass  $m_{\tilde{\chi}_1^0} > 46$  GeV, CL = 95%

[all  $\tan\beta$ , all  $m_0$ , all  $m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ ]

Mass  $m_{\tilde{\chi}_2^0} > 62.4$  GeV, CL = 95%

[ $1 < \tan\beta < 40$ , all  $m_0$ , all  $m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ ]

Mass  $m_{\tilde{\chi}_2^0} > 345$  GeV, CL = 95%

[ $\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W \tilde{\chi}_1^0 Z \tilde{\chi}_1^0$ , simplified model,  $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ ,  $m_{\tilde{\chi}_1^0} = 0$  GeV]

Mass  $m_{\tilde{\chi}_3^0} > 99.9$  GeV, CL = 95%

[ $1 < \tan\beta < 40$ , all  $m_0$ , all  $m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ ]

Mass  $m_{\tilde{\chi}_4^0} > 116$  GeV, CL = 95%

[ $1 < \tan\beta < 40$ , all  $m_0$ , all  $m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ ]

$\tilde{\chi}_i^\pm$  — charginos (mixtures of  $\tilde{W}^\pm$  and  $\tilde{H}_i^\pm$ )

Mass  $m_{\tilde{\chi}_1^\pm} > 94$  GeV, CL = 95%

[ $\tan\beta < 40$ ,  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} > 3$  GeV, all  $m_0$ ]

Mass  $m_{\tilde{\chi}_1^\pm} > 345$  GeV, CL = 95%

[simplified model,  $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ ,  $m_{\tilde{\chi}_1^0} = 0$  GeV]

$\tilde{\nu}$  — sneutrino

Mass  $m > 94$  GeV, CL = 95%

[CMSSM,  $1 \leq \tan\beta \leq 40$ ,  $m_{\tilde{e}_R} - m_{\tilde{\chi}_1^0} > 10$  GeV]

$\tilde{e}$  — scalar electron (selectron)

Mass  $m(\tilde{e}_L) > 107$  GeV, CL = 95% [all  $m_{\tilde{e}_R} - m_{\tilde{\chi}_1^0}$ ]

Mass  $m(\tilde{e}_R) > 97.5$  GeV, CL = 95%

[ $\Delta m > 11$  GeV,  $|\mu| > 100$  GeV,  $\tan\beta = 1.5$ ]

$\tilde{\mu}$  — scalar muon (smuon)

Mass  $m > 94$  GeV, CL = 95%

[CMSSM,  $1 \leq \tan\beta \leq 40$ ,  $m_{\tilde{\mu}_R} - m_{\tilde{\chi}_1^0} > 10$  GeV]

$\tilde{\tau}$  — scalar tau (stau)

Mass  $m > 81.9$  GeV, CL = 95%

[ $m_{\tilde{\tau}_R} - m_{\tilde{\chi}_1^0} > 15$  GeV, all  $\theta_\tau$ ,  $B(\tilde{\tau} \rightarrow \tau \tilde{\chi}_1^0) = 100\%$ ]

$\tilde{q}$  — squarks of the first two quark generations

The first of these limits is within CMSSM with cascade decays, evaluated assuming a fixed value of the parameters  $\mu$  and  $\tan\beta$ . The first two limits assume two-generations of mass degenerate squarks ( $\tilde{q}_L$  and  $\tilde{q}_R$ ) and gaugino mass parameters that are constrained by the unification condition at the grand unification scale. The third limit assumes a simplified model with a 100% branching ratio for the prompt decay  $\tilde{q} \rightarrow q \tilde{\chi}_1^0$ .

Mass  $m > 1450$  GeV, CL = 95%

[CMSSM,  $\tan\beta = 30$ ,  $A_0 = -2\max(m_0, m_{1/2})$ ,  $\mu > 0$ ]

Mass  $m > 850$  GeV, CL = 95%

[jets +  $\cancel{E}_T$ ,  $\tilde{q} \rightarrow q \tilde{\chi}_1^0$  simplified model,  $m_{\tilde{\chi}_1^0} = 0$  GeV]

Mass  $m > 520$  GeV, CL = 95%

[ $\tilde{q} \rightarrow q \tilde{\chi}_1^0$ , simplified model, single light squark,  $m_{\tilde{\chi}_1^0} = 0$ ]

$\tilde{b}$  — scalar bottom (sbottom)

Mass  $m > 650$  GeV, CL = 95% [ $\tilde{b} \rightarrow b \tilde{\chi}_1^0$ ,  $m_{\tilde{\chi}_1^0} = 0$ ]

Mass  $m > 600$  GeV, CL = 95% [ $\tilde{b} \rightarrow b \tilde{\chi}_1^0$ ,  $m_{\tilde{\chi}_1^0} < 250$  GeV]

$\tilde{t}$  — scalar top (stop)

Mass  $m > 730$  GeV, CL = 95%

[ $\tilde{t} \rightarrow t \tilde{\chi}_1^0$ ,  $m_{\tilde{\chi}_1^0} = 100$  GeV,  $m_{\tilde{t}} > m_t + m_{\tilde{\chi}_1^0}$ ]

Mass  $m > 500$  GeV, CL = 95%

[ $\ell^\pm + \text{jets} + \cancel{E}_T$ ,  $\tilde{t}_1 \rightarrow b \tilde{\chi}_1^\pm$ ,  $m_{\tilde{\chi}_1^\pm} = 2 m_{\tilde{\chi}_1^0}$ ,  $100 \text{ GeV} < m_{\tilde{\chi}_1^0} < 150$  GeV]

Mass  $m > 240$  GeV, CL = 95%

[ $\tilde{t}_1 \rightarrow c \tilde{\chi}_1^0$ ,  $m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} < 85$  GeV]

$\tilde{g}$  — gluino

The first limit assumes a simplified model with a 100% branching ratio for the prompt 3 body decay, independent of the squark mass. The second of these limits is within the CMSSM (for  $m_{\tilde{g}} \gtrsim 5$  GeV), and includes the effects of cascade decays, evaluated assuming a fixed value of the parameters  $\mu$  and  $\tan\beta$ . The limit assumes GUT relations between gaugino masses and the gauge couplings. The third limit is based on a combination of searches.

Mass  $m > 1225$  GeV, CL = 95% [ $\tilde{g} \rightarrow q \bar{q} \tilde{\chi}_1^0$ ,  $m_{\tilde{\chi}_1^0} = 0$ ]

Mass  $m > 1150$  GeV, CL = 95%

[CMSSM,  $\tan\beta = 30$ ,  $A_0 = -2\max(m_0, m_{1/2})$ ,  $\mu > 0$ ]

Mass  $m > 1150$  GeV, CL = 95%

[general RPC  $\tilde{g}$  decays,  $m_{\tilde{\chi}_1^0} < 100$  GeV]

## Technicolor

The limits for technicolor (and top-color) particles are quite varied depending on assumptions. See the Technicolor section of the full Review (the data listings).

## Quark and Lepton Compositeness, Searches for

### Scale Limits $\Lambda$ for Contact Interactions (the lowest dimensional interactions with four fermions)

If the Lagrangian has the form

$$\pm \frac{g^2}{2\Lambda^2} \bar{\psi}_L \gamma_\mu \psi_L \bar{\psi}_L \gamma^\mu \psi_L$$

(with  $g^2/4\pi$  set equal to 1), then we define  $\Lambda \equiv \Lambda_{LL}^\pm$ . For the full definitions and for other forms, see the Note in the Listings on Searches for Quark and Lepton Compositeness in the full Review and the original literature.

$$\Lambda_{LL}^+(eeee) > 8.3 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^-(eeee) > 10.3 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^+(ee\mu\mu) > 8.5 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^-(ee\mu\mu) > 9.5 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^+(ee\tau\tau) > 7.9 \text{ TeV, CL} = 95\%$$

# Searches Summary Table

$\Lambda_{LL}^-(e e \tau \tau)$	$> 7.2$ TeV, CL = 95%
$\Lambda_{LL}^+(\ell \ell \ell \ell)$	$> 9.1$ TeV, CL = 95%
$\Lambda_{LL}^-(\ell \ell \ell \ell)$	$> 10.3$ TeV, CL = 95%
$\Lambda_{LL}^+(e e u u)$	$> 23.3$ TeV, CL = 95%
$\Lambda_{LL}^-(e e u u)$	$> 12.5$ TeV, CL = 95%
$\Lambda_{LL}^+(e e d d)$	$> 11.1$ TeV, CL = 95%
$\Lambda_{LL}^-(e e d d)$	$> 26.4$ TeV, CL = 95%
$\Lambda_{LL}^+(e e c c)$	$> 9.4$ TeV, CL = 95%
$\Lambda_{LL}^-(e e c c)$	$> 5.6$ TeV, CL = 95%
$\Lambda_{LL}^+(e e b b)$	$> 9.4$ TeV, CL = 95%
$\Lambda_{LL}^-(e e b b)$	$> 10.2$ TeV, CL = 95%
$\Lambda_{LL}^+(\mu \mu q q)$	$> 12.5$ TeV, CL = 95%
$\Lambda_{LL}^-(\mu \mu q q)$	$> 16.7$ TeV, CL = 95%
$\Lambda(\ell \nu \ell \nu)$	$> 3.10$ TeV, CL = 90%
$\Lambda(e \nu q q)$	$> 2.81$ TeV, CL = 95%
$\Lambda_{LL}^+(q q q q)$	$> 9.0$ TeV, CL = 95%
$\Lambda_{LL}^-(q q q q)$	$> 12.0$ TeV, CL = 95%
$\Lambda_{LL}^+(\nu \nu q q)$	$> 5.0$ TeV, CL = 95%
$\Lambda_{LL}^-(\nu \nu q q)$	$> 5.4$ TeV, CL = 95%

## Excited Leptons

The limits from  $\ell^{*+} \ell^{*-}$  do not depend on  $\lambda$  (where  $\lambda$  is the  $\ell \ell^*$  transition coupling). The  $\lambda$ -dependent limits assume chiral coupling.

$e^{*\pm}$  — excited electron

Mass $m$	$> 103.2$ GeV, CL = 95% (from $e^* e^*$ )
Mass $m$	$> 3.000 \times 10^3$ GeV, CL = 95% (from $e e^*$ )
Mass $m$	$> 356$ GeV, CL = 95% (if $\lambda_\gamma = 1$ )

$\mu^{*\pm}$  — excited muon

Mass $m$	$> 103.2$ GeV, CL = 95% (from $\mu^* \mu^*$ )
Mass $m$	$> 3.000 \times 10^3$ GeV, CL = 95% (from $\mu \mu^*$ )

$\tau^{*\pm}$  — excited tau

Mass $m$	$> 103.2$ GeV, CL = 95% (from $\tau^* \tau^*$ )
Mass $m$	$> 2.500 \times 10^3$ GeV, CL = 95% (from $\tau \tau^*$ )

$\nu^*$  — excited neutrino

Mass $m$	$> 1.600 \times 10^3$ GeV, CL = 95% (from $\nu^* \nu^*$ )
Mass $m$	$> 213$ GeV, CL = 95% (from $\nu^* X$ )

$q^*$  — excited quark

Mass $m$	$> 338$ GeV, CL = 95% (from $q^* q^*$ )
Mass $m$	$> 4.060 \times 10^3$ GeV, CL = 95% (from $q^* X$ )

## Color Sextet and Octet Particles

Color Sextet Quarks ( $q_6$ )

Mass $m$	$> 84$ GeV, CL = 95% (Stable $q_6$ )
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Color Octet Charged Leptons ( $\ell_8$ )

Mass $m$	$> 86$ GeV, CL = 95% (Stable $\ell_8$ )
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Color Octet Neutrinos ( $\nu_8$ )

Mass $m$	$> 110$ GeV, CL = 90% ( $\nu_8 \rightarrow \nu g$ )
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## Extra Dimensions

Please refer to the Extra Dimensions section of the full *Review* for a discussion of the model-dependence of these bounds, and further constraints.

### Constraints on the radius of the extra dimensions, for the case of two-flat dimensions of equal radii

$R$	$< 30 \mu\text{m}$ , CL = 95% (direct tests of Newton's law)
$R$	$< 15 \mu\text{m}$ , CL = 95% ( $pp \rightarrow j G$ )
$R$	$< 0.16\text{--}916$ nm (astrophysics; limits depend on technique and assumptions)

### Constraints on the fundamental gravity scale

$M_{TT}$	$> 6.3$ TeV, CL = 95% ( $pp \rightarrow$ dijet, angular distribution)
$M_c$	$> 4.16$ TeV, CL = 95% ( $pp \rightarrow \ell \bar{\ell}$ )

### Constraints on the Kaluza-Klein graviton in warped extra dimensions

$M_G$	$> 2.73$ TeV, CL = 95% ( $pp \rightarrow e^+ e^-, \mu^+ \mu^-$ )
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### Constraints on the Kaluza-Klein gluon in warped extra dimensions

$M_{g_{KK}}$	$> 2.5$ TeV, CL = 95% ( $g_{KK} \rightarrow t \bar{t}$ )
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## TESTS OF CONSERVATION LAWS

Updated June 2016 by L. Wolfenstein (Carnegie-Mellon University) and C.-J. Lin (LBNL).

In keeping with the current interest in tests of conservation laws, we collect together a Table of experimental limits on all weak and electromagnetic decays, mass differences, and moments, and on a few reactions, whose observation would violate conservation laws. The Table is given only in the full *Review of Particle Physics*, not in the Particle Physics Booklet. For the benefit of Booklet readers, we include the best limits from the Table in the following text. Limits in this text are for CL=90% unless otherwise specified. The Table is in two parts: “Discrete Space-Time Symmetries,” *i.e.*,  $C$ ,  $P$ ,  $T$ ,  $CP$ , and  $CPT$ ; and “Number Conservation Laws,” *i.e.*, lepton, baryon, hadronic flavor, and charge conservation. The references for these data can be found in the the Particle Listings in the *Review*. A discussion of these tests follows.

### $CPT$ INVARIANCE

General principles of relativistic field theory require invariance under the combined transformation  $CPT$ . The simplest tests of  $CPT$  invariance are the equality of the masses and lifetimes of a particle and its antiparticle. The best test comes from the limit on the mass difference between  $K^0$  and  $\bar{K}^0$ . Any such difference contributes to the  $CP$ -violating parameter  $\epsilon$ . Assuming  $CPT$  invariance,  $\phi_\epsilon$ , the phase of  $\epsilon$  should be very close to  $44^\circ$ . (See the review “ $CP$  Violation in  $K_L$  decay” in this edition.) In contrast, if the entire source of  $CP$  violation in  $K^0$  decays were a  $K^0 - \bar{K}^0$  mass difference,  $\phi_\epsilon$  would be  $44^\circ + 90^\circ$ .

Assuming that there is no other source of  $CPT$  violation than this mass difference, it is possible to deduce that [1]

$$m_{\bar{K}^0} - m_{K^0} \approx \frac{2(m_{K_L^0} - m_{K_S^0}) |\eta| (\frac{2}{3}\phi_{+-} + \frac{1}{3}\phi_{00} - \phi_{SW})}{\sin \phi_{SW}},$$

where  $\phi_{SW} = (43.51 \pm 0.05)^\circ$ , the superweak angle. Using our best values of the  $CP$ -violation parameters, we get  $|(m_{\bar{K}^0} - m_{K^0})/m_{K^0}| \leq 0.6 \times 10^{-18}$  at CL=90%. Limits can also be placed on specific  $CPT$ -violating decay amplitudes. Given the small value of  $(1 - |\eta_{00}/\eta_{+-}|)$ , the value of  $\phi_{00} - \phi_{+-}$  provides a measure of  $CPT$  violation in  $K_L^0 \rightarrow 2\pi$  decay. Results from CERN [1] and Fermilab [2] indicate no  $CPT$ -violating effect.

### $CP$ AND $T$ INVARIANCE

Given  $CPT$  invariance,  $CP$  violation and  $T$  violation are equivalent. The original evidence for  $CP$  violation came from the measurement of  $|\eta_{+-}| = |A(K_L^0 \rightarrow \pi^+\pi^-)/A(K_S^0 \rightarrow \pi^+\pi^-)| = (2.232 \pm 0.011) \times 10^{-3}$ . This could be explained in terms of  $K^0 - \bar{K}^0$  mixing, which also leads to the asymmetry  $[\Gamma(K_L^0 \rightarrow \pi^- e^+ \nu) - \Gamma(K_L^0 \rightarrow \pi^+ e^- \bar{\nu})]/[\text{sum}] = (0.334 \pm 0.007)\%$ . Evidence for  $CP$  violation in the kaon decay amplitude comes from the measurement of  $(1 - |\eta_{00}/\eta_{+-}|)/3 = \text{Re}(\epsilon'/\epsilon) = (1.66 \pm 0.23) \times 10^{-3}$ . In the Standard Model much larger  $CP$ -violating effects are expected. The first of these, which is associated with  $B - \bar{B}$  mixing, is the parameter  $\sin(2\beta)$  now measured

quite accurately to be  $0.679 \pm 0.020$ . A number of other  $CP$ -violating observables are being measured in  $B$  decays; direct evidence for  $CP$  violation in the  $B$  decay amplitude comes from the asymmetry  $[\Gamma(\bar{B}^0 \rightarrow K^-\pi^+) - \Gamma(B^0 \rightarrow K^+\pi^-)]/[\text{sum}] = -0.082 \pm 0.006$ . Direct tests of  $T$  violation are much more difficult; a measurement by CPLEAR of the difference between the oscillation probabilities of  $K^0$  to  $\bar{K}^0$  and  $\bar{K}^0$  to  $K^0$  is related to  $T$  violation [3]. A nonzero value of the electric dipole moment of the neutron and electron requires both  $P$  and  $T$  violation. The current experimental results are  $< 3.0 \times 10^{-26}$  e cm (neutron), and  $< 8.7 \times 10^{-29}$  e cm (electron) at the 90% C.L. The BABAR experiment reported the first direct observation of  $T$  violation in the  $B$  system. The measured  $T$ -violating parameters in the time evolution of the neutral  $B$  mesons are  $\Delta S_T^+ = -1.37 \pm 0.15$  and  $\Delta S_T^- = 1.17 \pm 0.21$ , with a significance of  $14\sigma$  [4]. This observation of  $T$  violation, with exchange of initial and final states of the neutral  $B$ , was made possible in a  $B$ -factory using the Einstein-Podolsky-Rosen Entanglement of the two  $B$ 's produced in the decay of the  $\Upsilon(4S)$  and the two time-ordered decays of the  $B$ 's as filtering measurements of the meson state [5].

### CONSERVATION OF LEPTON NUMBERS

Present experimental evidence and the standard electroweak theory are consistent with the absolute conservation of three separate lepton numbers: electron number  $L_e$ , muon number  $L_\mu$ , and tau number  $L_\tau$ , except for the effect of neutrino mixing associated with neutrino masses. Searches for violations are of the following types:

**a)  $\Delta L = 2$  for one type of charged lepton.** The best limit comes from the search for neutrinoless double beta decay  $(Z, A) \rightarrow (Z+2, A) + e^- + e^-$ . The best laboratory limit is  $t_{1/2} > 1.07 \times 10^{26}$  yr (CL=90%) for  $^{136}\text{Xe}$  from the KamLAND-Zen experiment [6].

**b) Conversion of one charged-lepton type to another.** For purely leptonic processes, the best limits are on  $\mu \rightarrow e\gamma$  and  $\mu \rightarrow 3e$ , measured as  $\Gamma(\mu \rightarrow e\gamma)/\Gamma(\mu \rightarrow \text{all}) < 5.7 \times 10^{-13}$  and  $\Gamma(\mu \rightarrow 3e)/\Gamma(\mu \rightarrow \text{all}) < 1.0 \times 10^{-12}$ . For semileptonic processes, the best limit comes from the coherent conversion process in a muonic atom,  $\mu^- + (Z, A) \rightarrow e^- + (Z, A)$ , measured as  $\Gamma(\mu^- \text{Ti} \rightarrow e^- \text{Ti})/\Gamma(\mu^- \text{Ti} \rightarrow \text{all}) < 4.3 \times 10^{-12}$ . Of special interest is the case in which the hadronic flavor also changes, as in  $K_L \rightarrow e\mu$  and  $K^+ \rightarrow \pi^+ e^- \mu^+$ , measured as  $\Gamma(K_L \rightarrow e\mu)/\Gamma(K_L \rightarrow \text{all}) < 4.7 \times 10^{-12}$  and  $\Gamma(K^+ \rightarrow \pi^+ e^- \mu^+)/\Gamma(K^+ \rightarrow \text{all}) < 1.3 \times 10^{-11}$ . Limits on the conversion of  $\tau$  into  $e$  or  $\mu$  are found in  $\tau$  decay and are much less stringent than those for  $\mu \rightarrow e$  conversion, *e.g.*,  $\Gamma(\tau \rightarrow \mu\gamma)/\Gamma(\tau \rightarrow \text{all}) < 4.4 \times 10^{-8}$  and  $\Gamma(\tau \rightarrow e\gamma)/\Gamma(\tau \rightarrow \text{all}) < 3.3 \times 10^{-8}$ .

**c) Conversion of one type of charged lepton into another type of charged antilepton.** The case most studied is  $\mu^- + (Z, A) \rightarrow e^+ + (Z-2, A)$ , the strongest limit being  $\Gamma(\mu^- \text{Ti} \rightarrow e^+ \text{Ca})/\Gamma(\mu^- \text{Ti} \rightarrow \text{all}) < 3.6 \times 10^{-11}$ .

## Tests of Conservation Laws

**d) Neutrino oscillations.** It is expected even in the standard electroweak theory that the lepton numbers are not separately conserved, as a consequence of lepton mixing analogous to Cabibbo-Kobayashi-Maskawa quark mixing. However, if the only source of lepton-number violation is the mixing of low-mass neutrinos then processes such as  $\mu \rightarrow e\gamma$  are expected to have extremely small unobservable probabilities. For small neutrino masses, the lepton-number violation would be observed first in neutrino oscillations, which have been the subject of extensive experimental studies. Compelling evidence for neutrino mixing has come from atmospheric, solar, accelerator, and reactor neutrinos. Recently, the reactor neutrino experiments have measured the last neutrino mixing angle  $\theta_{13}$  and found it to be relatively large. For a comprehensive review on neutrino mixing, including the latest results on  $\theta_{13}$ , see the review “*Neutrino Mass, Mixing, and Oscillations*” by K. Nakamura and S.T. Petcov in this edition of RPP.

### CONSERVATION OF HADRONIC FLAVORS

In strong and electromagnetic interactions, hadronic flavor is conserved, *i.e.* the conversion of a quark of one flavor ( $d, u, s, c, b, t$ ) into a quark of another flavor is forbidden. In the Standard Model, the weak interactions violate these conservation laws in a manner described by the Cabibbo-Kobayashi-Maskawa mixing (see the section “Cabibbo-Kobayashi-Maskawa Mixing Matrix”). The way in which these conservation laws are violated is tested as follows:

**(a)  $\Delta S = \Delta Q$  rule.** In the strangeness-changing semileptonic decay of strange particles, the strangeness change equals the change in charge of the hadrons. Tests come from limits on decay rates such as  $\Gamma(\Sigma^+ \rightarrow ne^+\nu)/\Gamma(\Sigma^+ \rightarrow \text{all}) < 5 \times 10^{-6}$ , and from a detailed analysis of  $K_L \rightarrow \pi e \nu$ , which yields the parameter  $x$ , measured to be  $(\text{Re } x, \text{Im } x) = (-0.002 \pm 0.006, 0.0012 \pm 0.0021)$ . Corresponding rules are  $\Delta C = \Delta Q$  and  $\Delta B = \Delta Q$ .

**(b) Change of flavor by two units.** In the Standard Model this occurs only in second-order weak interactions. The classic example is  $\Delta S = 2$  via  $K^0 - \bar{K}^0$  mixing, which is directly measured by  $m(K_L) - m(K_S) = (0.5293 \pm 0.0009) \times 10^{10} \text{ } \hbar s^{-1}$ . The  $\Delta B = 2$  transitions in the  $B^0$  and  $B_s^0$  systems via mixing are also well established. The measured mass differences between the eigenstates are  $(m_{B_H^0} - m_{B_L^0}) = (0.5096 \pm 0.0034) \times 10^{12} \text{ } \hbar s^{-1}$  and  $(m_{B_{sH}^0} - m_{B_{sL}^0}) = (17.757 \pm 0.021) \times 10^{12} \text{ } \hbar s^{-1}$ . There is now strong evidence of  $\Delta C = 2$  transition in the charm sector with the mass difference  $m_{D_H^0} - m_{D_L^0} = (0.95^{+0.41}_{-0.44}) \times 10^{10} \text{ } \hbar s^{-1}$ . All results are consistent with the second-order calculations in the Standard Model.

**(c) Flavor-changing neutral currents.** In the Standard Model the neutral-current interactions do not change flavor. The low rate  $\Gamma(K_L \rightarrow \mu^+\mu^-)/\Gamma(K_L \rightarrow \text{all}) = (6.84 \pm 0.11) \times 10^{-9}$  puts limits on such interactions; the

nonzero value for this rate is attributed to a combination of the weak and electromagnetic interactions. The best test should come from  $K^+ \rightarrow \pi^+\nu\bar{\nu}$ , which occurs in the Standard Model only as a second-order weak process with a branching fraction of  $(0.4 \text{ to } 1.2) \times 10^{-10}$ . Combining results from BNL-E787 and BNL-E949 experiments yield  $\Gamma(K^+ \rightarrow \pi^+\nu\bar{\nu})/\Gamma(K^+ \rightarrow \text{all}) = (1.7 \pm 1.1) \times 10^{-10}$  [7]. Limits for charm-changing or bottom-changing neutral currents are less stringent:  $\Gamma(D^0 \rightarrow \mu^+\mu^-)/\Gamma(D^0 \rightarrow \text{all}) < 6.2 \times 10^{-9}$  and  $\Gamma(B^0 \rightarrow \mu^+\mu^-)/\Gamma(B^0 \rightarrow \text{all}) = (3.9^{+1.6}_{-1.4}) \times 10^{-10}$ . One cannot isolate flavor-changing neutral current (FCNC) effects in non leptonic decays. For example, the FCNC transition  $s \rightarrow d + (\bar{u} + u)$  is equivalent to the charged-current transition  $s \rightarrow u + (\bar{u} + d)$ . Tests for FCNC are therefore limited to hadron decays into lepton pairs. Such decays are expected only in second-order in the electroweak coupling in the Standard Model. The LHCb and CMS experiments have recently observed the FCNC decay of  $B_s^0 \rightarrow \mu^+\mu^-$ . The current world average value is  $\Gamma(B_s^0 \rightarrow \mu^+\mu^-)/\Gamma(B_s^0 \rightarrow \text{all}) = (2.9^{+0.7}_{-0.6}) \times 10^{-9}$ , which is consistent with the Standard Model expectation.

### References

1. R. Carosi *et al.*, Phys. Lett. **B237**, 303 (1990).
2. E. Abouzaid *et al.*, Phys. Rev. **D83**, 092001 (2011); B. Schwingerheuer *et al.*, Phys. Rev. Lett. **74**, 4376 (1995).
3. A. Angelopoulos *et al.*, Phys. Lett. **B444**, 43 (1998); L. Wolfenstein, Phys. Rev. Lett. **83**, 911 (1999).
4. J.P. Lees *et al.*, Phys. Rev. Lett. **109**, 211801 (2012).
5. M.C. Banuls, J. Bernabeu, Phys. Lett. **B464**, 117 (1999); Nucl. Phys. **B590**, 19 (2000).
6. A. Gando *et al.*, arXiv:1605.02889.
7. A.V. Artamonov *et al.*, Phys. Rev. Lett. **101**, 191802 (2008).

## TESTS OF DISCRETE SPACE-TIME SYMMETRIES

### CHARGE CONJUGATION (C) INVARIANCE

$\Gamma(\pi^0 \rightarrow 3\gamma)/\Gamma_{\text{total}}$	$< 3.1 \times 10^{-8}$ , CL = 90%
$\eta$ C-nonconserving decay parameters	
$\pi^+\pi^-\pi^0$ left-right asymmetry	$(0.09^{+0.11}_{-0.12}) \times 10^{-2}$
$\pi^+\pi^-\pi^0$ sextant asymmetry	$(0.12^{+0.10}_{-0.11}) \times 10^{-2}$
$\pi^+\pi^-\pi^0$ quadrant asymmetry	$(-0.09 \pm 0.09) \times 10^{-2}$
$\pi^+\pi^-\gamma$ left-right asymmetry	$(0.9 \pm 0.4) \times 10^{-2}$
$\pi^+\pi^-\gamma$ parameter $\beta$ (D-wave)	$-0.02 \pm 0.07$ (S = 1.3)
$\Gamma(\eta \rightarrow \pi^0\gamma)/\Gamma_{\text{total}}$	$< 9 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta \rightarrow 2\pi^0\gamma)/\Gamma_{\text{total}}$	$< 5 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta \rightarrow 3\pi^0\gamma)/\Gamma_{\text{total}}$	$< 6 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta \rightarrow 3\gamma)/\Gamma_{\text{total}}$	$< 1.6 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	[a] $< 4 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta \rightarrow \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	[a] $< 5 \times 10^{-6}$ , CL = 90%
$\Gamma(\omega(782) \rightarrow \eta\pi^0)/\Gamma_{\text{total}}$	$< 2.1 \times 10^{-4}$ , CL = 90%
$\Gamma(\omega(782) \rightarrow 2\pi^0)/\Gamma_{\text{total}}$	$< 2.1 \times 10^{-4}$ , CL = 90%
$\Gamma(\omega(782) \rightarrow 3\pi^0)/\Gamma_{\text{total}}$	$< 2.3 \times 10^{-4}$ , CL = 90%
asymmetry parameter for $\eta'(958) \rightarrow \pi^+\pi^-\gamma$ decay	$-0.03 \pm 0.04$
$\Gamma(\eta'(958) \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	[a] $< 1.4 \times 10^{-3}$ , CL = 90%

# Tests of Conservation Laws

$\Gamma(\eta'(958) \rightarrow \eta e^+ e^-)/\Gamma_{\text{total}}$	[a] $<2.4 \times 10^{-3}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow 3\gamma)/\Gamma_{\text{total}}$	$<1.0 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow \mu^+ \mu^- \pi^0)/\Gamma_{\text{total}}$	[a] $<6.0 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow \mu^+ \mu^- \eta)/\Gamma_{\text{total}}$	[a] $<1.5 \times 10^{-5}$ , CL = 90%
$\Gamma(J/\psi(1S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-7}$ , CL = 90%
$\Gamma(J/\psi(1S) \rightarrow \gamma\phi)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-6}$ , CL = 90%

## PARITY (P) INVARIANCE

$e$ electric dipole moment	$<0.87 \times 10^{-28}$ e cm, CL = 90%
$\mu$ electric dipole moment	$(-0.1 \pm 0.9) \times 10^{-19}$ e cm
$\text{Re}(d_\tau = \tau \text{ electric dipole moment})$	$-0.220 \text{ to } 0.45 \times 10^{-16}$ e cm, CL = 95%
$\Gamma(\eta \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta \rightarrow 2\pi^0)/\Gamma_{\text{total}}$	$<3.5 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta \rightarrow 4\pi^0)/\Gamma_{\text{total}}$	$<6.9 \times 10^{-7}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow \pi^0 \pi^0)/\Gamma_{\text{total}}$	$<4 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta_C(1S) \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$	$<1.1 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta_C(1S) \rightarrow \pi^0 \pi^0)/\Gamma_{\text{total}}$	$<4 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta_C(1S) \rightarrow K^+ K^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta_C(1S) \rightarrow K_S^0 \bar{K}_S^0)/\Gamma_{\text{total}}$	$<3.1 \times 10^{-4}$ , CL = 90%
$p$ electric dipole moment	$<0.54 \times 10^{-23}$ e cm
$n$ electric dipole moment	$<0.30 \times 10^{-25}$ e cm, CL = 90%
$A$ electric dipole moment	$<1.5 \times 10^{-16}$ e cm, CL = 95%

## TIME REVERSAL (T) INVARIANCE

$e$ electric dipole moment	$<0.87 \times 10^{-28}$ e cm, CL = 90%
$\mu$ electric dipole moment	$(-0.1 \pm 0.9) \times 10^{-19}$ e cm
$\mu$ decay parameters	
transverse $e^+$ polarization normal to plane of $\mu$ spin, $e^+$ momentum	$(-2 \pm 8) \times 10^{-3}$
$\alpha'/A$	$(-10 \pm 20) \times 10^{-3}$
$\beta'/A$	$(2 \pm 7) \times 10^{-3}$
$\text{Re}(d_\tau = \tau \text{ electric dipole moment})$	$-0.220 \text{ to } 0.45 \times 10^{-16}$ e cm, CL = 95%
$P_T$ in $K^+ \rightarrow \pi^0 \mu^+ \nu_\mu$	$(-1.7 \pm 2.5) \times 10^{-3}$
$P_T$ in $K^+ \rightarrow \mu^+ \nu_\mu \gamma$	$(-0.6 \pm 1.9) \times 10^{-2}$
$\text{Im}(\xi)$ in $K^+ \rightarrow \pi^0 \mu^+ \nu_\mu$ decay (from transverse $\mu$ pol.)	$-0.006 \pm 0.008$
asymmetry $A_T$ in $K^0-\bar{K}^0$ mixing	$(6.6 \pm 1.6) \times 10^{-3}$
$\text{Im}(\xi)$ in $K_{\mu 3}^0$ decay (from transverse $\mu$ pol.)	$-0.007 \pm 0.026$
$A_T(D^\pm \rightarrow K_S^0 K^\pm \pi^\mp)$	[b] $(-12 \pm 11) \times 10^{-3}$
$A_T(D^0 \rightarrow K^+ K^- \pi^+ \pi^-)$	[b] $(1.7 \pm 2.7) \times 10^{-3}$
$A_T(D_S^\pm \rightarrow K_S^0 K^\pm \pi^\mp)$	[b] $(-14 \pm 8) \times 10^{-3}$
$\Delta S_T^+(S_{\ell^-}^-, K_S^0 - S_{\ell^+}^+, K_S^0)$	$-1.37 \pm 0.15$
$\Delta S_T^-(S_{\ell^-}^+, K_S^0 - S_{\ell^+}^-, K_S^0)$	$1.17 \pm 0.21$
$\Delta C_T^+(C_{\ell^-}^-, K_S^0 - C_{\ell^+}^+, K_S^0)$	$0.10 \pm 0.16$
$\Delta C_T^-(C_{\ell^-}^+, K_S^0 - C_{\ell^+}^-, K_S^0)$	$0.04 \pm 0.16$
$p$ electric dipole moment	$<0.54 \times 10^{-23}$ e cm
$n$ electric dipole moment	$<0.30 \times 10^{-25}$ e cm, CL = 90%
$n \rightarrow p e^- \bar{\nu}_e$ decay parameters	
$\phi_{AV}$ , phase of $g_A$ relative to $g_V$	[c] $(180.017 \pm 0.026)^\circ$
triple correlation coefficient $D$	[d] $(-1.2 \pm 2.0) \times 10^{-4}$
triple correlation coefficient $R$	[d] $0.004 \pm 0.013$
$A$ electric dipole moment	$<1.5 \times 10^{-16}$ e cm, CL = 95%
triple correlation coefficient $D$ for $\Sigma^- \rightarrow n e^- \bar{\nu}_e$	$0.11 \pm 0.10$

## CP INVARIANCE

$\text{Re}(d_\tau^W)$	$<0.50 \times 10^{-17}$ e cm, CL = 95%
$\text{Im}(d_\tau^W)$	$<1.1 \times 10^{-17}$ e cm, CL = 95%
$\eta \rightarrow \pi^+ \pi^- e^+ e^-$ decay-plane asymmetry	$(-0.6 \pm 3.1) \times 10^{-2}$
$\Gamma(\eta \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta \rightarrow 2\pi^0)/\Gamma_{\text{total}}$	$<3.5 \times 10^{-4}$ , CL = 90%
$\Gamma(\eta \rightarrow 4\pi^0)/\Gamma_{\text{total}}$	$<6.9 \times 10^{-7}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-5}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow \pi^0 \pi^0)/\Gamma_{\text{total}}$	$<4 \times 10^{-4}$ , CL = 90%
$K^\pm \rightarrow \pi^\pm e^+ e^-$ rate difference/sum	$(-2.2 \pm 1.6) \times 10^{-2}$
$K^\pm \rightarrow \pi^\pm \mu^+ \mu^-$ rate difference/sum	$0.010 \pm 0.023$
$K^\pm \rightarrow \pi^\pm \pi^0 \gamma$ rate difference/sum	$(0.0 \pm 1.2) \times 10^{-3}$
$K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$ rate difference/sum	$(0.04 \pm 0.06)\%$
$K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$ rate difference/sum	$(-0.02 \pm 0.28)\%$
$K^\pm \rightarrow \pi^\pm \pi^+ \pi^- (g_+ - g_-) / (g_+ + g_-)$	$(-1.5 \pm 2.2) \times 10^{-4}$
$K^\pm \rightarrow \pi^\pm \pi^0 \pi^0 (g_+ - g_-) / (g_+ + g_-)$	$(1.8 \pm 1.8) \times 10^{-4}$
$A_S = [\Gamma(K_S^0 \rightarrow \pi^- e^+ \nu_e) - \Gamma(K_S^0 \rightarrow \pi^+ e^- \bar{\nu}_e)] / \text{SUM}$	$(2 \pm 10) \times 10^{-3}$
$\text{Im}(\eta_{+-0}) = \text{Im}(A(K_S^0 \rightarrow \pi^+ \pi^- \pi^0, CP\text{-violating}) / A(K_L^0 \rightarrow \pi^+ \pi^- \pi^0))$	$-0.002 \pm 0.009$
$\text{Im}(\eta_{000}) = \text{Im}(A(K_S^0 \rightarrow \pi^0 \pi^0 \pi^0) / A(K_L^0 \rightarrow \pi^0 \pi^0 \pi^0))$	$-0.001 \pm 0.016$
$ \eta_{000}  =  A(K_S^0 \rightarrow 3\pi^0) / A(K_L^0 \rightarrow 3\pi^0) $	$<0.0088$ , CL = 90%
$CP$ asymmetry $A$ in $K_S^0 \rightarrow \pi^+ \pi^- e^+ e^-$	$(-0.4 \pm 0.8)\%$
$\Gamma(K_S^0 \rightarrow 3\pi^0)/\Gamma_{\text{total}}$	$<2.6 \times 10^{-8}$ , CL = 90%
linear coefficient $j$ for $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$	$0.0012 \pm 0.0008$
quadratic coefficient $f$ for $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$	$0.004 \pm 0.006$
$ \epsilon'_{+-\gamma} /\epsilon$ for $K_L^0 \rightarrow \pi^+ \pi^- \gamma$	$<0.3$ , CL = 90%
$ g_{E1} $ for $K_L^0 \rightarrow \pi^+ \pi^- \gamma$	$<0.21$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	[e] $<3.8 \times 10^{-10}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	[e] $<2.8 \times 10^{-10}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	[f] $<2.6 \times 10^{-8}$ , CL = 90%
$A_{CP}(D^\pm \rightarrow \mu^\pm \nu)$	$(8 \pm 8)\%$
$A_{CP}(D^\pm \rightarrow K_L^0 e^\pm \nu)$	$(-0.6 \pm 1.6)\%$
$A_{CP}(D^\pm \rightarrow K_S^0 \pi^\pm)$	$(-0.41 \pm 0.09)\%$
$A_{CP}(D^\pm \rightarrow K^\mp 2\pi^\pm)$	$(-0.18 \pm 0.16)\%$
$A_{CP}(D^\pm \rightarrow K^\mp \pi^\pm \pi^\pm \pi^0)$	$(-0.3 \pm 0.7)\%$
$A_{CP}(D^\pm \rightarrow K_S^0 \pi^\pm \pi^0)$	$(-0.1 \pm 0.7)\%$
$A_{CP}(D^\pm \rightarrow K_S^0 \pi^\pm \pi^\mp \pi^-)$	$(0.0 \pm 1.2)\%$
$A_{CP}(D^\pm \rightarrow \pi^\pm \pi^0)$	$(2.9 \pm 2.9)\%$
$A_{CP}(D^\pm \rightarrow \pi^\pm \eta)$	$(1.0 \pm 1.5)\%$ ( $S = 1.4$ )
$A_{CP}(D^\pm \rightarrow \pi^\pm \eta'(958))$	$(-0.5 \pm 1.2)\%$ ( $S = 1.1$ )
$A_{CP}(\bar{K}^0 / K^0 K^\pm)$	$(0.11 \pm 0.17)\%$
$A_{CP}(D^\pm \rightarrow K_S^0 K^\pm)$	$(-0.11 \pm 0.25)\%$
$A_{CP}(D^\pm \rightarrow K^+ K^- \pi^\pm)$	$(0.37 \pm 0.29)\%$
$A_{CP}(D^\pm \rightarrow K^\pm K^* 0)$	$(-0.3 \pm 0.4)\%$
$A_{CP}(D^\pm \rightarrow \phi \pi^\pm)$	$(0.09 \pm 0.19)\%$ ( $S = 1.2$ )
$A_{CP}(D^\pm \rightarrow K^\pm K_0^*(1430)^0)$	$(8 \pm 7)\%$
$A_{CP}(D^\pm \rightarrow K^\pm K_2^*(1430)^0)$	$(43 \pm 20)\%$
$A_{CP}(D^\pm \rightarrow K^\pm K_0^*(800))$	$(-12 \pm 18)\%$
$A_{CP}(D^\pm \rightarrow a_0(1450)^0 \pi^\pm)$	$(-19 \pm 14)\%$
$A_{CP}(D^\pm \rightarrow \phi(1680) \pi^\pm)$	$(-9 \pm 26)\%$
$A_{CP}(D^\pm \rightarrow \pi^+ \pi^- \pi^\pm)$	$(-2 \pm 4)\%$
$A_{CP}(D^\pm \rightarrow K_S^0 K^\pm \pi^\mp \pi^-)$	$(-4 \pm 7)\%$
$A_{CP}(D^\pm \rightarrow K^\pm \pi^0)$	$(-4 \pm 11)\%$
Local $CPV$ in $D^\pm \rightarrow \pi^+ \pi^- \pi^\pm$	78.1%
Local $CPV$ in $D^\pm \rightarrow K^+ K^- \pi^\pm$	31%
$ q/p $ of $D^0-\bar{D}^0$ mixing	$0.92 \pm 0.12$
$A_\Gamma$ of $D^0-\bar{D}^0$ mixing	$(-0.125 \pm 0.526) \times 10^{-3}$
Where there is ambiguity, the $CP$ test is labelled by the $D^0$ decay mode.	
$A_{CP}(D^0 \rightarrow K^+ K^-)$	$(-0.14 \pm 0.12)\%$
$A_{CP}(D^0 \rightarrow K_S^0 K_S^0)$	$(-5 \pm 5)\%$
$A_{CP}(D^0 \rightarrow \pi^+ \pi^-)$	$(0.01 \pm 0.15)\%$
$A_{CP}(D^0 \rightarrow \pi^0 \pi^0)$	$(0.0 \pm 0.6)\%$
$A_{CP}(D^0 \rightarrow \pi^+ \pi^- \pi^0)$	$(0.3 \pm 0.4)\%$
$A_{CP}(D^0 \rightarrow \rho(770)^+ \pi^- \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(1.2 \pm 0.9)\%$
$A_{CP}(D^0 \rightarrow \rho(770)^0 \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(-3.1 \pm 3.0)\%$

# Tests of Conservation Laws

$ACP(D^0 \rightarrow \rho(770)^- \pi^+ \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(-1.0 \pm 1.7)\%$	$ACP(D^0 \rightarrow K_1^+(1270)^- K^+ \rightarrow \rho^0 K^- K^+)$	$(10 \pm 13)\%$
$ACP(D^0 \rightarrow \rho(1450)^+ \pi^- \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(0 \pm 70)\%$	$ACP(D^0 \rightarrow K^*(1410)^+ K^- \rightarrow K^{*0} \pi^+ K^-)$	$(-20 \pm 17)\%$
$ACP(D^0 \rightarrow \rho(1450)^0 \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(-20 \pm 40)\%$	$ACP(D^0 \rightarrow K^*(1410)^- K^+ \rightarrow \overline{K}^{*0} \pi^- K^+)$	$(-1 \pm 14)\%$
$ACP(D^0 \rightarrow \rho(1450)^- \pi^+ \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(6 \pm 9)\%$	$ACP(D^0 \rightarrow K^{*0} \overline{K}^{*0} S\text{-wave})$	$(10 \pm 14)\%$
$ACP(D^0 \rightarrow \rho(1700)^+ \pi^- \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(-5 \pm 14)\%$	$ACP(D^0 \rightarrow \phi \rho^0 S\text{-wave})$	$(-3 \pm 5)\%$
$ACP(D^0 \rightarrow \rho(1700)^0 \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(13 \pm 9)\%$	$ACP(D^0 \rightarrow \phi \rho^0 D\text{-wave})$	$(-37 \pm 19)\%$
$ACP(D^0 \rightarrow \rho(1700)^- \pi^+ \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(8 \pm 11)\%$	$ACP(D^0 \rightarrow \phi(\pi^+ \pi^-) S\text{-wave})$	$(-9 \pm 10)\%$
$ACP(D^0 \rightarrow f_0(980) \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(0 \pm 35)\%$	$ACP((K^- \pi^+) P\text{-wave} (K^+ \pi^-) S\text{-wave})$	$(3 \pm 11)\%$
$ACP(D^0 \rightarrow f_0(1370) \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(25 \pm 18)\%$	CP-even fraction in $D^0 \rightarrow \pi^+ \pi^- \pi^0$ decays	$(97.3 \pm 1.7)\%$
$ACP(D^0 \rightarrow f_0(1500) \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(0 \pm 18)\%$	CP-even fraction in $D^0 \rightarrow K^+ K^- \pi^0$ decays	$(73 \pm 6)\%$
$ACP(D^0 \rightarrow f_0(1710) \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(0 \pm 24)\%$	CP-even fraction in $D^0 \rightarrow \pi^+ \pi^- \pi^+ \pi^-$ decays	$(73.7 \pm 2.8)\%$
$ACP(D^0 \rightarrow f_2(1270) \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(-4 \pm 6)\%$	$\Delta A_{CP}^{D^0} = A_{CP}(K^+ K^-) - A_{CP}(\pi^+ \pi^-)$	$(-0.32 \pm 0.22)\%$ ( $S = 1.9$ )
$ACP(D^0 \rightarrow \sigma(400) \pi^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(6 \pm 8)\%$	Local CPV in $D^0, \overline{D}^0 \rightarrow \pi^+ \pi^- \pi^0$	4.9%
$ACP(\text{nonresonant } D^0 \rightarrow \pi^+ \pi^- \pi^0)$	[g] $(-13 \pm 23)\%$	Local CPV in $D^0, \overline{D}^0 \rightarrow \pi^+ \pi^- \pi^+ \pi^-$	41%
$ACP(D^0 \rightarrow K^+ K^- \pi^0)$	$(-1.0 \pm 1.7)\%$	Local CPV in $D^0, \overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$	96%
$ACP(D^0 \rightarrow K^*(892)^+ K^- \rightarrow K^+ K^- \pi^0)$	[g] $(-0.9 \pm 1.3)\%$	Local CPV in $D^0, \overline{D}^0 \rightarrow K^+ K^- \pi^0$	16.6%
$ACP(D^0 \rightarrow K^*(1410)^+ K^- \rightarrow K^+ K^- \pi^0)$	[g] $(-21 \pm 24)\%$	Local CPV in $D^0, \overline{D}^0 \rightarrow K^+ K^- \pi^+ \pi^-$	9.1%
$ACP(D^0 \rightarrow (K^+ \pi^0)_S K^- \rightarrow K^+ K^- \pi^0)$	[g] $(7 \pm 15)\%$	$ACP(D_S^{\pm\pm} \rightarrow \mu^{\pm\pm} \nu)$	$(5 \pm 6)\%$
$ACP(D^0 \rightarrow \phi(1020) \pi^0 \rightarrow K^+ K^- \pi^0)$	[g] $(1.1 \pm 2.2)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} K_S^0)$	$(0.08 \pm 0.26)\%$
$ACP(D^0 \rightarrow f_0(980) \pi^0 \rightarrow K^+ K^- \pi^0)$	[g] $(-3 \pm 19)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^+ K^- \pi^{\pm})$	$(-0.5 \pm 0.9)\%$
$ACP(D^0 \rightarrow a_0(980) \pi^0 \rightarrow K^+ K^- \pi^0)$	[g] $(-5 \pm 16)\%$	$ACP(D_S^{\pm\pm} \rightarrow \phi \pi^{\pm})$	$(-0.38 \pm 0.27)\%$
$ACP(D^0 \rightarrow f_2'(1525) \pi^0 \rightarrow K^+ K^- \pi^0)$	[g] $(0 \pm 160)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} K_S^0 \pi^0)$	$(-2 \pm 6)\%$
$ACP(D^0 \rightarrow K^*(892)^- K^+ \rightarrow K^+ K^- \pi^0)$	[g] $(-5 \pm 4)\%$	$ACP(D_S^{\pm\pm} \rightarrow 2K_S^0 \pi^{\pm})$	$(3 \pm 5)\%$
$ACP(D^0 \rightarrow K^*(1410)^- K^+ \rightarrow K^+ K^- \pi^0)$	[g] $(-17 \pm 29)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^+ K^- \pi^{\pm} \pi^0)$	$(0.0 \pm 3.0)\%$
$ACP(D^0 \rightarrow (K^- \pi^0)_S K^+ \rightarrow K^+ K^- \pi^0)$	[g] $(-10 \pm 40)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} K_S^0 \pi^+ \pi^-)$	$(-6 \pm 5)\%$
$ACP(D^0 \rightarrow K_S^0 \pi^0)$	$(-0.20 \pm 0.17)\%$	$ACP(D_S^{\pm\pm} \rightarrow K_S^0 K^{\mp} 2\pi^{\pm})$	$(4.1 \pm 2.8)\%$
$ACP(D^0 \rightarrow K_S^0 \eta)$	$(0.5 \pm 0.5)\%$	$ACP(D_S^{\pm\pm} \rightarrow \pi^+ \pi^- \pi^{\pm})$	$(-0.7 \pm 3.1)\%$
$ACP(D^0 \rightarrow K_S^0 \eta')$	$(1.0 \pm 0.7)\%$	$ACP(D_S^{\pm\pm} \rightarrow \pi^{\pm} \eta)$	$(1.1 \pm 3.1)\%$
$ACP(D^0 \rightarrow K_S^0 \phi)$	$(-3 \pm 9)\%$	$ACP(D_S^{\pm\pm} \rightarrow \pi^{\pm} \eta')$	$(-2.2 \pm 2.3)\%$
$ACP(D^0 \rightarrow K^- \pi^+)$	$(0.3 \pm 0.7)\%$	$ACP(D_S^{\pm\pm} \rightarrow \eta \pi^{\pm} \pi^0)$	$(-1 \pm 4)\%$
$ACP(D^0 \rightarrow K^+ \pi^-)$	$(0.0 \pm 1.6)\%$	$ACP(D_S^{\pm\pm} \rightarrow \eta' \pi^{\pm} \pi^0)$	$(0 \pm 8)\%$
$ACP(D^0 \rightarrow K^- \pi^+ \pi^0)$	$(0.1 \pm 0.5)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} \pi^0)$	$(-27 \pm 24)\%$
$ACP(D^0 \rightarrow K^+ \pi^- \pi^0)$	$(0 \pm 5)\%$	$ACP(\overline{K}^0 / K^0 \pi^{\pm})$	$(0.4 \pm 0.5)\%$
$ACP(D^0 \rightarrow K_S^0 \pi^+ \pi^-)$	$(-0.1 \pm 0.8)\%$	$ACP(D_S^{\pm\pm} \rightarrow K_S^0 \pi^{\pm})$	$(3.1 \pm 2.6)\%$ ( $S = 1.7$ )
$ACP(D^0 \rightarrow K^*(892)^- \pi^+ \rightarrow K_S^0 \pi^+ \pi^-)$	$(0.4 \pm 0.5)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} \pi^+ \pi^-)$	$(4 \pm 5)\%$
$ACP(D^0 \rightarrow K^*(892)^+ \pi^- \rightarrow K_S^0 \pi^+ \pi^-)$	$(1 \pm 6)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} \eta)$	$(9 \pm 15)\%$
$ACP(D^0 \rightarrow K_S^0 \rho^0 \rightarrow K_S^0 \pi^+ \pi^-)$	$(-0.1 \pm 0.5)\%$	$ACP(D_S^{\pm\pm} \rightarrow K^{\pm} \eta'(958))$	$(6 \pm 19)\%$
$ACP(D^0 \rightarrow K_S^0 \omega \rightarrow K_S^0 \pi^+ \pi^-)$	$(-13 \pm 7)\%$	$ACP(B^+ \rightarrow J/\psi(1S) K^+)$	$0.003 \pm 0.006$ ( $S = 1.8$ )
$ACP(D^0 \rightarrow K_S^0 f_0(980) \rightarrow K_S^0 \pi^+ \pi^-)$	$(-0.4 \pm 2.7)\%$	$ACP(B^+ \rightarrow J/\psi(1S) \pi^+)$	$(0.1 \pm 2.8) \times 10^{-2}$ ( $S = 1.2$ )
$ACP(D^0 \rightarrow K_S^0 f_2(1270) \rightarrow K_S^0 \pi^+ \pi^-)$	$(-4 \pm 5)\%$	$ACP(B^+ \rightarrow J/\psi \rho^+)$	$-0.11 \pm 0.14$
$ACP(D^0 \rightarrow K_S^0 f_0(1370) \rightarrow K_S^0 \pi^+ \pi^-)$	$(-1 \pm 9)\%$	$ACP(B^+ \rightarrow J/\psi K^*(892)^+)$	$-0.048 \pm 0.033$
$ACP(D^0 \rightarrow \overline{K}^0 \rho^0(1450) \rightarrow K_S^0 \pi^+ \pi^-)$	$(-4 \pm 10)\%$	$ACP(B^+ \rightarrow \eta_c K^+)$	$0.01 \pm 0.07$ ( $S = 2.2$ )
$ACP(D^0 \rightarrow \overline{K}^0 f_0(600) \rightarrow K_S^0 \pi^+ \pi^-)$	$(-3 \pm 5)\%$	$ACP(B^+ \rightarrow \psi(2S) \pi^+)$	$0.03 \pm 0.06$
$ACP(D^0 \rightarrow K^*(1410)^- \pi^+ \rightarrow K_S^0 \pi^+ \pi^-)$	$(-2 \pm 9)\%$	$ACP(B^+ \rightarrow \psi(2S) K^+)$	$0.012 \pm 0.020$ ( $S = 1.5$ )
$ACP(D^0 \rightarrow K_S^0 \pi^+ \pi^-)$	$(4 \pm 4)\%$	$ACP(B^+ \rightarrow \psi(2S) K^*(892)^+)$	$0.08 \pm 0.21$
$ACP(D^0 \rightarrow K_S^0(1430)^- \pi^+ \rightarrow K_S^0 \pi^+ \pi^-)$	$(12 \pm 15)\%$	$ACP(B^+ \rightarrow \chi_{c1}(1P) \pi^+)$	$0.07 \pm 0.18$
$ACP(D^0 \rightarrow K_S^0(1430)^- \pi^+ \rightarrow K_S^0 \pi^+ \pi^-)$	$(3 \pm 6)\%$	$ACP(B^+ \rightarrow \chi_{c0} K^+)$	$-0.20 \pm 0.18$ ( $S = 1.5$ )
$ACP(D^0 \rightarrow K_2^*(1430)^+ \pi^- \rightarrow K_S^0 \pi^+ \pi^-)$	$(-10 \pm 32)\%$	$ACP(B^+ \rightarrow \chi_{c1} K^+)$	$-0.009 \pm 0.033$
$ACP(D^0 \rightarrow K^*(1680)^- \pi^+ \rightarrow K_S^0 \pi^+ \pi^-)$	—	$ACP(B^+ \rightarrow \chi_{c1} K^*(892)^+)$	$0.5 \pm 0.5$
$ACP(D^0 \rightarrow K^- \pi^+ \pi^+ \pi^-)$	$(0.2 \pm 0.5)\%$	$ACP(B^+ \rightarrow \overline{D}^0 \pi^+)$	$-0.007 \pm 0.007$
$ACP(D^0 \rightarrow K^+ \pi^- \pi^+ \pi^-)$	$(-2 \pm 4)\%$	$ACP(B^+ \rightarrow D_{CP(+1)} \pi^+)$	$0.035 \pm 0.024$
$ACP(D^0 \rightarrow K^+ K^- \pi^+ \pi^-)$	$(-8 \pm 7)\%$	$ACP(B^+ \rightarrow D_{CP(-1)} \pi^+)$	$0.017 \pm 0.026$
$ACP(D^0 \rightarrow K_1^+(1270)^+ K^- \rightarrow K^{*0} \pi^+ K^-)$	$(-1 \pm 10)\%$	$ACP(B^+ \rightarrow \overline{D}^0 K^+)$	$0.007 \pm 0.025$ ( $S = 1.5$ )
$ACP(D^0 \rightarrow K_1^+(1270)^- K^+ \rightarrow \overline{K}^{*0} \pi^- K^+)$	$(-10 \pm 32)\%$	$r_B(B^+ \rightarrow D^0 K^+)$	$0.095 \pm 0.008$
$ACP(D^0 \rightarrow K_1^+(1270)^+ K^- \rightarrow \rho^0 K^+ K^-)$	$(-7 \pm 17)\%$	$\delta_B(B^+ \rightarrow D^0 K^+)$	$(123 \pm 10)^\circ$
		$r_B(B^+ \rightarrow \overline{D}^0 K^{*+})$	$0.17 \pm 0.11$ ( $S = 2.3$ )
		$\delta_B(B^+ \rightarrow D^0 K^{*+})$	$(155 \pm 70)^\circ$ ( $S = 2.0$ )
		$ACP(B^+ \rightarrow [K^- \pi^+]_D K^+)$	$-0.58 \pm 0.21$
		$ACP(B^+ \rightarrow [K^- \pi^+ \pi^0]_D K^+)$	$0.07 \pm 0.30$ ( $S = 1.5$ )
		$ACP(B^+ \rightarrow [K^+ K^- \pi^0]_D K^+)$	$0.30 \pm 0.20$
		$ACP(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D K^+)$	$0.05 \pm 0.09$
		$ACP(B^+ \rightarrow [K^- \pi^+]_{\overline{D}} K^*(892)^+)$	$-0.3 \pm 0.5$
		$ACP(B^+ \rightarrow [K^- \pi^+]_D \pi^+)$	$0.00 \pm 0.09$
		$ACP(B^+ \rightarrow [K^- \pi^+ \pi^0]_D \pi^+)$	$0.35 \pm 0.16$

Unless otherwise stated, limits are given at the 90% confidence level, while errors are given as  $\pm 1$  standard deviation.

## Tests of Conservation Laws

$A_{CP}(B^+ \rightarrow [K^+ K^- \pi^0]_D \pi^+)$	$-0.03 \pm 0.04$	$A_{CP}(B^+ \rightarrow \rho^0 K^*(892)^+)$	$0.31 \pm 0.13$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D \pi^+)$	$-0.016 \pm 0.020$	$A_{CP}(B^+ \rightarrow K^*(892)^+ f_0(980))$	$-0.15 \pm 0.12$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\pi)} \pi^+)$	$-0.09 \pm 0.27$	$A_{CP}(B^+ \rightarrow a_1^+ K^0)$	$0.12 \pm 0.11$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\gamma)} \pi^+)$	$-0.7 \pm 0.6$	$A_{CP}(B^+ \rightarrow b_1^+ K^0)$	$-0.03 \pm 0.15$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\pi)} K^+)$	$0.8 \pm 0.4$	$A_{CP}(B^+ \rightarrow K^*(892)^0 \rho^+)$	$-0.01 \pm 0.16$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\gamma)} K^+)$	$0.4 \pm 1.0$	$A_{CP}(B^+ \rightarrow b_1^0 K^+)$	$-0.46 \pm 0.20$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D K^+)$	$-0.02 \pm 0.15$	$A_{CP}(B^+ \rightarrow K^0 K^+)$	$0.04 \pm 0.14$
$A_{CP}(B^+ \rightarrow [K_S^0 K^+ \pi^-]_D K^+)$	$0.04 \pm 0.09$	$A_{CP}(B^+ \rightarrow K^+ K_S^0 K_S^0)$	$0.04^{+0.04}_{-0.05}$
$A_{CP}(B^+ \rightarrow [K_S^0 K^- \pi^+]_D K^+)$	$0.23 \pm 0.13$	$A_{CP}(B^+ \rightarrow K^+ K^- \pi^+)$	$-0.118 \pm 0.022$
$A_{CP}(B^+ \rightarrow [K_S^0 K^- \pi^+]_D \pi^+)$	$-0.052 \pm 0.034$	$A_{CP}(B^+ \rightarrow K^+ K^- K^+)$	$-0.033 \pm 0.008$
$A_{CP}(B^+ \rightarrow [K_S^0 K^+ \pi^-]_D \pi^+)$	$-0.025 \pm 0.026$	$A_{CP}(B^+ \rightarrow \phi K^+)$	$0.024 \pm 0.028$ ( $S = 2.3$ )
$A_{CP}(B^+ \rightarrow [K^*(892)^- K^+]_D K^+)$	$0.03 \pm 0.11$	$A_{CP}(B^+ \rightarrow X_0(1550) K^+)$	$-0.04 \pm 0.07$
$A_{CP}(B^+ \rightarrow [K^*(892)^+ K^-]_D K^+)$	$0.34 \pm 0.21$	$A_{CP}(B^+ \rightarrow K^{*+} K^+ K^-)$	$0.11 \pm 0.09$
$A_{CP}(B^+ \rightarrow [K^*(892)^+ K^-]_D \pi^+)$	$-0.05 \pm 0.05$	$A_{CP}(B^+ \rightarrow \phi K^*(892)^+)$	$-0.01 \pm 0.08$
$A_{CP}(B^+ \rightarrow [K^*(892)^- K^+]_D \pi^+)$	$-0.012 \pm 0.030$	$A_{CP}(B^+ \rightarrow \phi(K\pi)_0^{*+})$	$0.04 \pm 0.16$
$A_{CP}(B^+ \rightarrow D_{CP(+1)} K^+)$	$0.170 \pm 0.033$ ( $S = 1.2$ )	$A_{CP}(B^+ \rightarrow \phi K_1(1270)^+)$	$0.15 \pm 0.20$
$A_{ADS}(B^+ \rightarrow D K^+)$	$-0.52 \pm 0.15$	$A_{CP}(B^+ \rightarrow \phi K_2^*(1430)^+)$	$-0.23 \pm 0.20$
$A_{ADS}(B^+ \rightarrow D \pi^+)$	$0.14 \pm 0.06$	$A_{CP}(B^+ \rightarrow K^+ \phi \phi)$	$-0.10 \pm 0.08$
$A_{ADS}(B^+ \rightarrow [K^- \pi^+]_D K^+ \pi^- \pi^+)$	$-0.33 \pm 0.35$	$A_{CP}(B^+ \rightarrow K^+ [\phi \phi]_{\eta_c})$	$0.09 \pm 0.10$
$A_{ADS}(B^+ \rightarrow [K^- \pi^+]_D \pi^+ \pi^- \pi^+)$	$-0.01 \pm 0.09$	$A_{CP}(B^+ \rightarrow K^*(892)^+ \gamma)$	$0.018 \pm 0.029$
$A_{CP}(B^+ \rightarrow D_{CP(-1)} K^+)$	$-0.10 \pm 0.07$	$A_{CP}(B^+ \rightarrow \eta K^+ \gamma)$	$-0.12 \pm 0.07$
$A_{CP}(B^+ \rightarrow [K^+ K^-]_D K^+ \pi^- \pi^+)$	$-0.04 \pm 0.06$	$A_{CP}(B^+ \rightarrow \phi K^+ \gamma)$	$-0.13 \pm 0.11$ ( $S = 1.1$ )
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^-]_D K^+ \pi^- \pi^+)$	$-0.05 \pm 0.10$	$A_{CP}(B^+ \rightarrow \rho^+ \gamma)$	$-0.11 \pm 0.33$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_D K^+ \pi^- \pi^+)$	$0.013 \pm 0.023$	$A_{CP}(B^+ \rightarrow \pi^+ \pi^0)$	$0.03 \pm 0.04$
$A_{CP}(B^+ \rightarrow [K^+ K^-]_D \pi^+ \pi^- \pi^+)$	$-0.019 \pm 0.015$	$A_{CP}(B^+ \rightarrow \pi^+ \pi^- \pi^+)$	$0.057 \pm 0.013$
$A_{CP}(B^+ \rightarrow [\pi^+ \pi^-]_D \pi^+ \pi^- \pi^+)$	$-0.013 \pm 0.019$	$A_{CP}(B^+ \rightarrow \rho^0 \pi^+)$	$0.18^{+0.09}_{-0.17}$
$A_{CP}(B^+ \rightarrow [K^- \pi^+]_D \pi^+ \pi^- \pi^+)$	$-0.002 \pm 0.011$	$A_{CP}(B^+ \rightarrow f_2(1270) \pi^+)$	$0.41 \pm 0.30$
$A_{CP}(B^+ \rightarrow \bar{D}^{*0} \pi^+)$	$-0.014 \pm 0.015$	$A_{CP}(B^+ \rightarrow \rho^0(1450) \pi^+)$	$-0.1^{+0.4}_{-0.5}$
$A_{CP}(B^+ \rightarrow (D_{CP(+1)}^*)^0 \pi^+)$	$-0.02 \pm 0.05$	$A_{CP}(B^+ \rightarrow \pi^+ \pi^- \pi^+ \text{ nonresonant})$	$-0.14^{+0.23}_{-0.16}$
$A_{CP}(B^+ \rightarrow (D_{CP(-1)}^*)^0 \pi^+)$	$-0.09 \pm 0.05$	$A_{CP}(B^+ \rightarrow \rho^+ \pi^0)$	$0.02 \pm 0.11$
$A_{CP}(B^+ \rightarrow D^{*0} K^+)$	$-0.07 \pm 0.04$	$A_{CP}(B^+ \rightarrow \rho^+ \rho^0)$	$-0.05 \pm 0.05$
$r_B^*(B^+ \rightarrow D^{*0} K^+)$	$0.114^{+0.023}_{-0.040}$ ( $S = 1.2$ )	$A_{CP}(B^+ \rightarrow \omega \pi^+)$	$-0.04 \pm 0.06$
$\delta_B^*(B^+ \rightarrow D^{*0} K^+)$	$(310^{+22}_{-26})^\circ$ ( $S = 1.3$ )	$A_{CP}(B^+ \rightarrow \omega \rho^+)$	$-0.20 \pm 0.09$
$A_{CP}(B^+ \rightarrow D_{CP(+1)}^{*0} K^+)$	$-0.12 \pm 0.08$	$A_{CP}(B^+ \rightarrow \eta \pi^+)$	$-0.14 \pm 0.07$ ( $S = 1.4$ )
$A_{CP}(B^+ \rightarrow D_{CP(-1)}^{*0} K^+)$	$0.07 \pm 0.10$	$A_{CP}(B^+ \rightarrow \eta \rho^+)$	$0.11 \pm 0.11$
$A_{CP}(B^+ \rightarrow D_{CP(+1)} K^* (892)^+$	$0.09 \pm 0.14$	$A_{CP}(B^+ \rightarrow \eta' \pi^+)$	$0.06 \pm 0.16$
$A_{CP}(B^+ \rightarrow D_{CP(-1)} K^* (892)^+$	$-0.23 \pm 0.22$	$A_{CP}(B^+ \rightarrow \eta' \rho^+)$	$0.26 \pm 0.17$
$A_{CP}(B^+ \rightarrow D_S^+ \phi)$	$0.0 \pm 0.4$	$A_{CP}(B^+ \rightarrow b_1^0 \pi^+)$	$0.05 \pm 0.16$
$A_{CP}(B^+ \rightarrow D^{*+} \bar{D}^{*0})$	$-0.15 \pm 0.11$	$A_{CP}(B^+ \rightarrow p \bar{p} \pi^+)$	$0.00 \pm 0.04$
$A_{CP}(B^+ \rightarrow D^{*+} \bar{D}^0)$	$-0.06 \pm 0.13$	$A_{CP}(B^+ \rightarrow p \bar{p} K^+)$	$0.00 \pm 0.04$ ( $S = 2.2$ )
$A_{CP}(B^+ \rightarrow D^+ \bar{D}^{*0})$	$0.13 \pm 0.18$	$A_{CP}(B^+ \rightarrow p \bar{p} K^*(892)^+)$	$0.21 \pm 0.16$ ( $S = 1.4$ )
$A_{CP}(B^+ \rightarrow D^+ \bar{D}^0)$	$-0.03 \pm 0.07$	$A_{CP}(B^+ \rightarrow p \bar{p} \gamma)$	$0.17 \pm 0.17$
$A_{CP}(B^+ \rightarrow K_S^0 \pi^+)$	$-0.017 \pm 0.016$	$A_{CP}(B^+ \rightarrow p \bar{p} \pi^0)$	$0.01 \pm 0.17$
$A_{CP}(B^+ \rightarrow K^+ \pi^0)$	$0.037 \pm 0.021$	$A_{CP}(B^+ \rightarrow K^+ \ell^+ \ell^-)$	$-0.02 \pm 0.08$
$A_{CP}(B^+ \rightarrow \eta' K^+)$	$0.004 \pm 0.011$	$A_{CP}(B^+ \rightarrow K^+ e^+ e^-)$	$0.14 \pm 0.14$
$A_{CP}(B^+ \rightarrow \eta' K^*(892)^+)$	$-0.26 \pm 0.27$	$A_{CP}(B^+ \rightarrow K^+ \mu^+ \mu^-)$	$0.011 \pm 0.017$
$A_{CP}(B^+ \rightarrow \eta' K_0^*(1430)^+)$	$0.06 \pm 0.20$	$A_{CP}(B^+ \rightarrow K^{*+} \ell^+ \ell^-)$	$-0.09 \pm 0.14$
$A_{CP}(B^+ \rightarrow \eta' K_2^*(1430)^+)$	$0.15 \pm 0.13$	$A_{CP}(B^+ \rightarrow K^{*+} e^+ e^-)$	$-0.14 \pm 0.23$
$A_{CP}(B^+ \rightarrow \eta K^*(892)^+)$	$0.02 \pm 0.06$	$A_{CP}(B^+ \rightarrow K^{*+} \mu^+ \mu^-)$	$-0.12 \pm 0.24$
$A_{CP}(B^+ \rightarrow \eta K_0^*(1430)^+)$	$0.05 \pm 0.13$	$\text{Re}(\epsilon_{B0})/(1+ \epsilon_{B0} ^2)$	$(-0.4 \pm 0.4) \times 10^{-3}$
$A_{CP}(B^+ \rightarrow \eta K_2^*(1430)^+)$	$-0.45 \pm 0.30$	$A_{T/CP}$	$0.005 \pm 0.018$
$A_{CP}(B^+ \rightarrow \omega K^+)$	$-0.02 \pm 0.04$	$A_{CP}(B^0 \rightarrow D^*(2010)^+ D^-)$	$0.037 \pm 0.034$
$A_{CP}(B^+ \rightarrow \omega K^{*+})$	$0.29 \pm 0.35$	$A_{CP}(B^0 \rightarrow [K^+ K^-]_D K^*(892)^0)$	$-0.20 \pm 0.15$
$A_{CP}(B^+ \rightarrow \omega(K\pi)_0^{*+})$	$-0.10 \pm 0.09$	$A_{CP}(B^0 \rightarrow [K^+ \pi^-]_D K^*(892)^0)$	$-0.03 \pm 0.04$
$A_{CP}(B^+ \rightarrow \omega K_2^*(1430)^+)$	$0.14 \pm 0.15$	$A_{CP}(B^0 \rightarrow [\pi^+ \pi^-]_D K^*(892)^0)$	$-0.09 \pm 0.22$
$A_{CP}(B^+ \rightarrow K^{*0} \pi^+)$	$-0.04 \pm 0.09$ ( $S = 2.1$ )	$A_{CP}(B^0 \rightarrow \eta' K^*(892)^0)$	$-0.07 \pm 0.18$
$A_{CP}(B^+ \rightarrow K^*(892)^+ \pi^0)$	$-0.06 \pm 0.24$	$A_{CP}(B^0 \rightarrow \eta' K_0^*(1430)^0)$	$-0.19 \pm 0.17$
$A_{CP}(B^+ \rightarrow K^+ \pi^- \pi^+)$	$0.027 \pm 0.008$	$A_{CP}(B^0 \rightarrow \eta' K_2^*(1430)^0)$	$0.14 \pm 0.18$
$A_{CP}(B^+ \rightarrow K^+ K^- K^+ \text{ nonresonant})$	$0.06 \pm 0.05$	$A_{CP}(B^0 \rightarrow \eta K_0^*(1430)^0)$	$0.06 \pm 0.13$
$A_{CP}(B^+ \rightarrow f(980)^0 K^+)$	$-0.08 \pm 0.09$	$A_{CP}(B^0 \rightarrow \eta K_2^*(1430)^0)$	$-0.07 \pm 0.19$
$A_{CP}(B^+ \rightarrow f_0(1500) K^+)$	$0.28 \pm 0.30$	$A_{CP}(B^0 \rightarrow b_1 K^+)$	$-0.07 \pm 0.12$
$A_{CP}(B^+ \rightarrow f'_2(1525)^0 K^+)$	$-0.08^{+0.05}_{-0.04}$	$A_{CP}(B^0 \rightarrow \omega K^{*0})$	$0.45 \pm 0.25$
$A_{CP}(B^+ \rightarrow K_0^*(1430)^0 \pi^+)$	$0.055 \pm 0.033$	$A_{CP}(B^0 \rightarrow \omega(K\pi)_0^0)$	$-0.07 \pm 0.09$
$A_{CP}(B^+ \rightarrow K_2^*(1430)^0 \pi^+)$	$0.05^{+0.29}_{-0.24}$	$A_{CP}(B^0 \rightarrow \omega K_2^*(1430)^0)$	$-0.37 \pm 0.17$
$A_{CP}(B^+ \rightarrow K^+ \pi^0 \pi^0)$	$-0.06 \pm 0.07$	$A_{CP}(B^0 \rightarrow K^+ \pi^- \pi^0)$	$(0 \pm 6) \times 10^{-2}$
$A_{CP}(B^+ \rightarrow K^0 \rho^+)$	$-0.12 \pm 0.17$	$A_{CP}(B^0 \rightarrow \rho^- K^+)$	$0.20 \pm 0.11$
$A_{CP}(B^+ \rightarrow K^{*+} \pi^+ \pi^-)$	$0.07 \pm 0.08$	$A_{CP}(B^0 \rightarrow \rho(1450)^- K^+)$	$-0.10 \pm 0.33$
		$A_{CP}(B^0 \rightarrow \rho(1700)^- K^+)$	$-0.4 \pm 0.6$
		$A_{CP}(B^0 \rightarrow K^+ \pi^- \pi^0 \text{ nonresonant})$	$0.10 \pm 0.18$

Unless otherwise stated, limits are given at the 90% confidence level, while errors are given as  $\pm 1$  standard deviation.

# Tests of Conservation Laws

$A_{CP}(B^0 \rightarrow K^0 \pi^+ \pi^-)$	$-0.01 \pm 0.05$	$C_{K+K-K_S^0}(B^0 \rightarrow K^+ K^- K_S^0)$	$0.06 \pm 0.08$
$A_{CP}(B^0 \rightarrow K^*(892)^+ \pi^-)$	$-0.22 \pm 0.06$	nonresonant)	
$A_{CP}(B^0 \rightarrow (K\pi)_0^{*+} \pi^-)$	$0.09 \pm 0.07$	$C_{K+K-K_S^0}(B^0 \rightarrow K^+ K^- K_S^0 \text{ inclusive})$	$0.01 \pm 0.09$
$A_{CP}(B^0 \rightarrow (K\pi)_0^{*0} \pi^0)$	$-0.15 \pm 0.11$	$C_{\phi K_S^0}(B^0 \rightarrow \phi K_S^0)$	$0.01 \pm 0.14$
$A_{CP}(B^0 \rightarrow K^{*0} \pi^0)$	$-0.15 \pm 0.13$	$S_{\phi K_S^0}(B^0 \rightarrow \phi K_S^0)$	$0.59 \pm 0.14$
$A_{CP}(B^0 \rightarrow K^*(892)^0 \pi^+ \pi^-)$	$0.07 \pm 0.05$	$C_{K_S K_S K_S}(B^0 \rightarrow K_S K_S K_S)$	$-0.23 \pm 0.14$
$A_{CP}(B^0 \rightarrow K^*(892)^0 \rho^0)$	$-0.06 \pm 0.09$	$S_{K_S K_S K_S}(B^0 \rightarrow K_S K_S K_S)$	$-0.5 \pm 0.6 \text{ (S = 3.0)}$
$A_{CP}(B^0 \rightarrow K^{*0} f_0(980))$	$0.07 \pm 0.10$	$C_{K_S^0 \pi^0 \gamma}(B^0 \rightarrow K_S^0 \pi^0 \gamma)$	$0.36 \pm 0.33$
$A_{CP}(B^0 \rightarrow K^{*+} \rho^-)$	$0.21 \pm 0.15$	$S_{K_S^0 \pi^0 \gamma}(B^0 \rightarrow K_S^0 \pi^0 \gamma)$	$-0.8 \pm 0.6$
$A_{CP}(B^0 \rightarrow K^*(892)^0 K^+ K^-)$	$0.01 \pm 0.05$	$C_{K^*(892)^0 \gamma}(B^0 \rightarrow K^*(892)^0 \gamma)$	$-0.04 \pm 0.16 \text{ (S = 1.2)}$
$A_{CP}(B^0 \rightarrow a_1^- K^+)$	$-0.16 \pm 0.12$	$S_{K^*(892)^0 \gamma}(B^0 \rightarrow K^*(892)^0 \gamma)$	$-0.15 \pm 0.22$
$A_{CP}(B^0 \rightarrow K^0 K^0)$	$-0.6 \pm 0.7$	$C_{\eta K^0 \gamma}(B^0 \rightarrow \eta K^0 \gamma)$	$-0.3 \pm 0.4$
$A_{CP}(B^0 \rightarrow K^*(892)^0 \phi)$	$0.00 \pm 0.04$	$S_{\eta K^0 \gamma}(B^0 \rightarrow \eta K^0 \gamma)$	$-0.2 \pm 0.5$
$A_{CP}(B^0 \rightarrow K^*(892)^0 K^- \pi^+)$	$0.2 \pm 0.4$	$C_{K^0 \phi \gamma}(B^0 \rightarrow K^0 \phi \gamma)$	$-0.3 \pm 0.6$
$A_{CP}(B^0 \rightarrow \phi(K\pi)_0^{*0})$	$0.12 \pm 0.08$	$S_{K^0 \phi \gamma}(B^0 \rightarrow K^0 \phi \gamma)$	$0.7^{+0.7}_{-1.1}$
$A_{CP}(B^0 \rightarrow \phi K_2^*(1430)^0)$	$-0.11 \pm 0.10$	$C(B^0 \rightarrow K_S^0 \rho^0 \gamma)$	$-0.05 \pm 0.19$
$A_{CP}(B^0 \rightarrow K^*(892)^0 \gamma)$	$-0.002 \pm 0.015$	$S(B^0 \rightarrow K_S^0 \rho^0 \gamma)$	$0.11 \pm 0.34$
$A_{CP}(B^0 \rightarrow K_2^*(1430)^0 \gamma)$	$-0.08 \pm 0.15$	$C(B^0 \rightarrow \rho^0 \gamma)$	$0.4 \pm 0.5$
$A_{CP}(B^0 \rightarrow \rho^+ \pi^-)$	$0.13 \pm 0.06 \text{ (S = 1.1)}$	$S(B^0 \rightarrow \rho^0 \gamma)$	$-0.8 \pm 0.7$
$A_{CP}(B^0 \rightarrow \rho^- \pi^+)$	$-0.08 \pm 0.08$	$C_{\pi\pi}(B^0 \rightarrow \pi^+ \pi^-)$	$-0.31 \pm 0.05$
$A_{CP}(B^0 \rightarrow a_1(1260)^\pm \pi^\mp)$	$-0.07 \pm 0.06$	$C_{\pi^0 \pi^0}(B^0 \rightarrow \pi^0 \pi^0)$	$-0.43 \pm 0.24$
$A_{CP}(B^0 \rightarrow b_1^- \pi^+)$	$-0.05 \pm 0.10$	$C_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-)$	$-0.03 \pm 0.07 \text{ (S = 1.2)}$
$A_{CP}(B^0 \rightarrow \rho \bar{\rho} K^*(892)^0)$	$0.05 \pm 0.12$	$S_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-)$	$0.05 \pm 0.07$
$A_{CP}(B^0 \rightarrow \rho \bar{\Lambda} \pi^-)$	$0.04 \pm 0.07$	$\Delta S_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-)$	$0.01 \pm 0.08$
$A_{CP}(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$	$-0.05 \pm 0.10$	$C_{\rho^0 \pi^0}(B^0 \rightarrow \rho^0 \pi^0)$	$0.27 \pm 0.24$
$A_{CP}(B^0 \rightarrow K^{*0} e^+ e^-)$	$-0.21 \pm 0.19$	$S_{\rho^0 \pi^0}(B^0 \rightarrow \rho^0 \pi^0)$	$-0.23 \pm 0.34$
$A_{CP}(B^0 \rightarrow K^{*0} \mu^+ \mu^-)$	$-0.034 \pm 0.024$	$C_{a_1\pi}(B^0 \rightarrow a_1(1260)^+ \pi^-)$	$-0.05 \pm 0.11$
$C_{D^*(2010)^- D^+}(B^0 \rightarrow D^*(2010)^- D^+)$	$-0.01 \pm 0.11$	$S_{a_1\pi}(B^0 \rightarrow a_1(1260)^+ \pi^-)$	$-0.2 \pm 0.4 \text{ (S = 3.2)}$
$C_{D^*(2010)^+ D^-}(B^0 \rightarrow D^*(2010)^+ D^-)$	$0.00 \pm 0.13 \text{ (S = 1.3)}$	$\Delta C_{a_1\pi}(B^0 \rightarrow a_1(1260)^+ \pi^-)$	$0.43 \pm 0.14 \text{ (S = 1.3)}$
$C_{D^{*+} D^{*-}}(B^0 \rightarrow D^{*+} D^{*-})$	$0.01 \pm 0.09 \text{ (S = 1.6)}$	$\Delta S_{a_1\pi}(B^0 \rightarrow a_1(1260)^+ \pi^-)$	$-0.11 \pm 0.12$
$C_+(B^0 \rightarrow D^{*+} D^{*-})$	$0.00 \pm 0.10 \text{ (S = 1.6)}$	$C(B^0 \rightarrow b_1^- K^+)$	$-0.22 \pm 0.24$
$C_-(B^0 \rightarrow D^{*+} D^{*-})$	$0.19 \pm 0.31$	$\Delta C(B^0 \rightarrow b_1^- \pi^+)$	$-1.04 \pm 0.24$
$S_-(B^0 \rightarrow D^{*+} D^{*-})$	$0.1 \pm 1.6 \text{ (S = 3.5)}$	$C_{\rho^0 \rho^0}(B^0 \rightarrow \rho^0 \rho^0)$	$0.2 \pm 0.9$
$C(B^0 \rightarrow D^*(2010)^+ D^*(2010)^- K_S^0)$	$0.01 \pm 0.29$	$S_{\rho^0 \rho^0}(B^0 \rightarrow \rho^0 \rho^0)$	$0.3 \pm 0.7$
$S(B^0 \rightarrow D^*(2010)^+ D^*(2010)^- K_S^0)$	$0.1 \pm 0.4$	$C_{\rho\rho}(B^0 \rightarrow \rho^+ \rho^-)$	$0.00 \pm 0.09$
$C_{D^+ D^-}(B^0 \rightarrow D^+ D^-)$	$-0.46 \pm 0.21 \text{ (S = 1.8)}$	$S_{\rho\rho}(B^0 \rightarrow \rho^+ \rho^-)$	$-0.14 \pm 0.13$
$C_{J/\psi(1S) \pi^0}(B^0 \rightarrow J/\psi(1S) \pi^0)$	$-0.13 \pm 0.13$	$ \lambda (B^0 \rightarrow J/\psi K^*(892)^0)$	$<0.25, \text{ CL} = 95\%$
$C(B^0 \rightarrow J/\psi(1S) \rho^0)$	$-0.06 \pm 0.06$	$\cos 2\beta(B^0 \rightarrow J/\psi K^*(892)^0)$	$1.7^{+0.7}_{-0.9} \text{ (S = 1.6)}$
$C_{D^{(*)} h^0}(B^0 \rightarrow D_{CP}^{(*)} h^0)$	$-0.02 \pm 0.08$	$\cos 2\beta(B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0)$	$1.0^{+0.6}_{-0.7} \text{ (S = 1.8)}$
$S_{D^{(*)} h^0}(B^0 \rightarrow D_{CP}^{(*)} h^0)$	$-0.66 \pm 0.12$	$(S_+ + S_-)/2(B^0 \rightarrow D^{*-} \pi^+)$	$-0.039 \pm 0.011$
$C_{K^0 \pi^0}(B^0 \rightarrow K^0 \pi^0)$	$0.00 \pm 0.13 \text{ (S = 1.4)}$	$(S_- - S_+)/2(B^0 \rightarrow D^{*-} \pi^+)$	$-0.009 \pm 0.015$
$C_{\eta'(958) K_S^0}(B^0 \rightarrow \eta'(958) K_S^0)$	$-0.04 \pm 0.20 \text{ (S = 2.5)}$	$(S_+ + S_-)/2(B^0 \rightarrow D^- \pi^+)$	$-0.046 \pm 0.023$
$S_{\eta'(958) K_S^0}(B^0 \rightarrow \eta'(958) K_S^0)$	$0.43 \pm 0.17 \text{ (S = 1.5)}$	$(S_- - S_+)/2(B^0 \rightarrow D^- \pi^+)$	$-0.022 \pm 0.021$
$C_{\eta' K^0}(B^0 \rightarrow \eta' K^0)$	$-0.06 \pm 0.04$	$(S_+ + S_-)/2(B^0 \rightarrow D^- \rho^+)$	$-0.024 \pm 0.032$
$C_{\omega K_S^0}(B^0 \rightarrow \omega K_S^0)$	$0.0 \pm 0.4 \text{ (S = 3.0)}$	$(S_- - S_+)/2(B^0 \rightarrow D^- \rho^+)$	$-0.10 \pm 0.06$
$S_{\omega K_S^0}(B^0 \rightarrow \omega K_S^0)$	$0.70 \pm 0.21$	$C_{\eta_c K_S^0}(B^0 \rightarrow \eta_c K_S^0)$	$0.08 \pm 0.13$
$C(B^0 \rightarrow K_S^0 \pi^0 \pi^0)$	$0.2 \pm 0.5$	$C_{c\bar{c}K^{(*)0}}(B^0 \rightarrow c\bar{c}K^{(*)0})$	$(0.5 \pm 1.7) \times 10^{-2}$
$S(B^0 \rightarrow K_S^0 \pi^0 \pi^0)$	$0.7 \pm 0.7$	$C_{J/\psi(nS) K^0}(B^0 \rightarrow J/\psi(nS) K^0)$	$(0.5 \pm 2.0) \times 10^{-2}$
$C_{\rho^0 K_S^0}(B^0 \rightarrow \rho^0 K_S^0)$	$-0.04 \pm 0.20$	$C_{J/\psi K^{*0}}(B^0 \rightarrow J/\psi K^{*0})$	$0.03 \pm 0.10$
$S_{\rho^0 K_S^0}(B^0 \rightarrow \rho^0 K_S^0)$	$0.50^{+0.17}_{-0.21}$	$S_{J/\psi K^{*0}}(B^0 \rightarrow J/\psi K^{*0})$	$0.60 \pm 0.25$
$C_{f_0(980) K_S^0}(B^0 \rightarrow f_0(980) K_S^0)$	$0.29 \pm 0.20$	$C_{\chi_{c0} K_S^0}(B^0 \rightarrow \chi_{c0} K_S^0)$	$-0.3^{+0.5}_{-0.4}$
$S_{f_0(980) K_S^0}(B^0 \rightarrow f_0(980) K_S^0)$	$-0.50 \pm 0.16$	$S_{\chi_{c0} K_S^0}(B^0 \rightarrow \chi_{c0} K_S^0)$	$-0.7 \pm 0.5$
$S_{f_2(1270) K_S^0}(B^0 \rightarrow f_2(1270) K_S^0)$	$-0.5 \pm 0.5$	$C_{\chi_{c1} K_S^0}(B^0 \rightarrow \chi_{c1} K_S^0)$	$0.06 \pm 0.07$
$C_{f_2(1270) K_S^0}(B^0 \rightarrow f_2(1270) K_S^0)$	$0.3 \pm 0.4$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow \phi K^0)$	$0.22 \pm 0.30$
$S_{f_x(1300) K_S^0}(B^0 \rightarrow f_x(1300) K_S^0)$	$-0.2 \pm 0.5$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow \phi K_0^*(1430)^0)$	$0.97^{+0.03}_{-0.52}$
$C_{f_x(1300) K_S^0}(B^0 \rightarrow f_x(1300) K_S^0)$	$0.13 \pm 0.35$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0)$	$0.45 \pm 0.28$
$S_{K^0 \pi^+ \pi^-}(B^0 \rightarrow K^0 \pi^+ \pi^- \text{ nonresonant})$	$-0.01 \pm 0.33$	$ \lambda (B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0)$	$1.01 \pm 0.08$
$C_{K^0 \pi^+ \pi^-}(B^0 \rightarrow K^0 \pi^+ \pi^- \text{ nonresonant})$	$0.01 \pm 0.26$	$ \sin(2\beta + \gamma) $	$>0.40, \text{ CL} = 90\%$
$C_{K_S^0 K_S^0}(B^0 \rightarrow K_S^0 K_S^0)$	$0.0 \pm 0.4 \text{ (S = 1.4)}$	$2\beta + \gamma$	$(83 \pm 60)^\circ$
$S_{K_S^0 K_S^0}(B^0 \rightarrow K_S^0 K_S^0)$	$-0.8 \pm 0.5$	$\gamma(B^0 \rightarrow D^0 K^{*0})$	$(162 \pm 60)^\circ$
		$A_{CP}(B \rightarrow K^*(892) \gamma)$	$-0.003 \pm 0.017$

Unless otherwise stated, limits are given at the 90% confidence level, while errors are given as  $\pm 1$  standard deviation.

# Tests of Conservation Laws

$ACP(b \rightarrow s\gamma)$	$0.015 \pm 0.020$	$ACP(B^0 \rightarrow K^+\pi^-)$	$-0.082 \pm 0.006$
$ACP(b \rightarrow (s+d)\gamma)$	$0.010 \pm 0.031$	$ACP(B^0 \rightarrow \eta K^*(892)^0)$	$0.19 \pm 0.05$
$ACP(B \rightarrow X_s \ell^+ \ell^-)$	$0.04 \pm 0.11$	$S_{D^*(2010)-D^+}(B^0 \rightarrow D^*(2010)^- D^+)$	$-0.72 \pm 0.15$
$ACP(B \rightarrow K^* e^+ e^-)$	$-0.18 \pm 0.15$	$S_{D^*(2010)+D^-}(B^0 \rightarrow D^*(2010)^+ D^-)$	$-0.73 \pm 0.14$
$ACP(B \rightarrow K^* \mu^+ \mu^-)$	$-0.03 \pm 0.13$	$S_{D^{*+}D^{*-}}(B^0 \rightarrow D^{*+} D^{*-})$	$-0.59 \pm 0.14$ (S = 1.8)
$ACP(B \rightarrow K^* \ell^+ \ell^-)$	$-0.04 \pm 0.07$	$S_+(B^0 \rightarrow D^{*+} D^{*-})$	$-0.73 \pm 0.09$
$ACP(B \rightarrow \eta \text{anything})$	$-0.13^{+0.04}_{-0.05}$	$S_{D^+D^-}(B^0 \rightarrow D^+ D^-)$	$-0.99^{+0.17}_{-0.14}$
$\text{Re}(\epsilon_{B_S^0}) / (1 +  \epsilon_{B_S^0} ^2)$	$(-1.9 \pm 1.0) \times 10^{-3}$	$S_{J/\psi(1S)\pi^0}(B^0 \rightarrow J/\psi(1S)\pi^0)$	$-0.94 \pm 0.29$ (S = 1.9)
CP Violation phase $\beta_s$	$(0.6 \pm 1.9) \times 10^{-2}$ rad	$S(B^0 \rightarrow J/\psi(1S)\rho^0)$	$-0.66^{+0.16}_{-0.12}$
$ACP(B_S \rightarrow \pi^+ K^-)$	$0.263 \pm 0.035$	$S_{K^0\pi^0}(B^0 \rightarrow K^0\pi^0)$	$0.58 \pm 0.17$
$ACP(B_S^0 \rightarrow [K^+ K^-]_D \overline{K}^*(892)^0)$	$-0.04 \pm 0.07$	$S_{\eta' K^0}(B^0 \rightarrow \eta' K^0)$	$0.63 \pm 0.06$
$\Gamma(\eta_c(1S) \rightarrow \pi^+ \pi^-) / \Gamma_{\text{total}}$	$< 1.1 \times 10^{-4}$ , CL = 90%	$S_{K^+ K^- K_S^0}(B^0 \rightarrow K^+ K^- K_S^0)$	$-0.66 \pm 0.11$
$\Gamma(\eta_c(1S) \rightarrow \pi^0 \pi^0) / \Gamma_{\text{total}}$	$< 4 \times 10^{-5}$ , CL = 90%	nonresonant)	
$\Gamma(\eta_c(1S) \rightarrow K^+ K^-) / \Gamma_{\text{total}}$	$< 6 \times 10^{-4}$ , CL = 90%	$S_{K^+ K^- K_S^0}(B^0 \rightarrow K^+ K^- K_S^0 \text{ inclusive})$	$-0.65 \pm 0.12$
$\Gamma(\eta_c(1S) \rightarrow K_S^0 K_S^0) / \Gamma_{\text{total}}$	$< 3.1 \times 10^{-4}$ , CL = 90%	$S_{\pi\pi}(B^0 \rightarrow \pi^+ \pi^-)$	$-0.67 \pm 0.06$
$(\alpha + \overline{\alpha}) / (\alpha - \overline{\alpha})$ in $\Lambda \rightarrow p\pi^-, \overline{\Lambda} \rightarrow \overline{p}\pi^+$	$0.006 \pm 0.021$	$\Delta C_{\rho\pi}(B^0 \rightarrow \rho^+ \pi^-)$	$0.27 \pm 0.06$
$\frac{[\alpha(\Xi^-)\alpha_-(\Lambda) - \alpha(\Xi^+)\alpha_+(\overline{\Lambda})]}{[\alpha(\Xi^-)\alpha_-(\Lambda) + \alpha(\Xi^+)\alpha_+(\overline{\Lambda})]}$	$(0 \pm 7) \times 10^{-4}$	$S_{\eta_c K_S^0}(B^0 \rightarrow \eta_c K_S^0)$	$0.93 \pm 0.17$
$(\alpha + \overline{\alpha}) / (\alpha - \overline{\alpha})$ in $\Omega^- \rightarrow \Lambda K^-, \overline{\Omega}^+ \rightarrow \overline{\Lambda} K^+$	$-0.02 \pm 0.13$	$\sin(2\beta)(B^0 \rightarrow J/\psi K_S^0)$	$0.679 \pm 0.020$
$(\alpha + \overline{\alpha}) / (\alpha - \overline{\alpha})$ in $\Lambda_c^+ \rightarrow \Lambda \pi^+, \overline{\Lambda}_c^- \rightarrow \overline{\Lambda} \pi^-$	$-0.07 \pm 0.31$	$S_{J/\psi(nS)K^0}(B^0 \rightarrow J/\psi(nS)K^0)$	$0.676 \pm 0.021$
$(\alpha + \overline{\alpha}) / (\alpha - \overline{\alpha})$ in $\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e, \overline{\Lambda}_c^- \rightarrow \overline{\Lambda} e^- \overline{\nu}_e$	$0.00 \pm 0.04$	$S_{\chi_{c1} K_S^0}(B^0 \rightarrow \chi_{c1} K_S^0)$	$0.63 \pm 0.10$
$ACP(\Lambda_b \rightarrow p\pi^-)$	$0.06 \pm 0.07$	$\sin(2\beta_{\text{eff}})(B^0 \rightarrow K^+ K^- K_S^0)$	$0.77^{+0.13}_{-0.12}$
$ACP(\Lambda_b \rightarrow pK^-)$	$0.00 \pm 0.19$ (S = 2.4)	$\alpha$	$(93 \pm 5)^\circ$
		$\text{Re}(\epsilon_b) / (1 +  \epsilon_b ^2)$	$(1.2 \pm 0.4) \times 10^{-3}$

## CP VIOLATION OBSERVED

$\text{Re}(\epsilon)$	$(1.596 \pm 0.013) \times 10^{-3}$
charge asymmetry in $K_{L3}^0$ decays	
$A_L = \text{weighted average of } A_L(\mu) \text{ and } A_L(e)$	$(0.332 \pm 0.006)\%$
$A_L(\mu) = [\Gamma(\pi^+ \mu^+ \nu_\mu) - \Gamma(\pi^+ \mu^- \overline{\nu}_\mu)] / \text{sum}$	$(0.304 \pm 0.025)\%$
$A_L(e) = [\Gamma(\pi^+ e^+ \nu_e) - \Gamma(\pi^+ e^- \overline{\nu}_e)] / \text{sum}$	$(0.334 \pm 0.007)\%$
parameters for $K_L^0 \rightarrow 2\pi$ decay	
$ \eta_{00}  =  A(K_L^0 \rightarrow 2\pi^0)  / A(K_S^0 \rightarrow 2\pi^0) $	$(2.220 \pm 0.011) \times 10^{-3}$ (S = 1.8)
$ \eta_{+-}  =  A(K_L^0 \rightarrow \pi^+ \pi^-)  / A(K_S^0 \rightarrow \pi^+ \pi^-) $	$(2.232 \pm 0.011) \times 10^{-3}$ (S = 1.8)
$ \epsilon  = (2 \eta_{+-}  +  \eta_{00} ) / 3$	$(2.228 \pm 0.011) \times 10^{-3}$ (S = 1.8)
$ \eta_{00} / \eta_{+-} $	$[h] \ 0.9950 \pm 0.0007$ (S = 1.6)
$\text{Re}(\epsilon' / \epsilon) = (1 -  \eta_{00} / \eta_{+-} ) / 3$	$[h] \ (1.66 \pm 0.23) \times 10^{-3}$ (S = 1.6)
Assuming CPT	
$\phi_{+-}$ , phase of $\eta_{+-}$	$(43.51 \pm 0.05)^\circ$ (S = 1.2)
$\phi_{00}$ , phase of $\eta_{00}$	$(43.52 \pm 0.05)^\circ$ (S = 1.3)
$\phi_\epsilon = (2\phi_{+-} + \phi_{00}) / 3$	$(43.52 \pm 0.05)^\circ$ (S = 1.2)
Not assuming CPT	
$\phi_{+-}$ , phase of $\eta_{+-}$	$(43.4 \pm 0.5)^\circ$ (S = 1.2)
$\phi_{00}$ , phase of $\eta_{00}$	$(43.7 \pm 0.6)^\circ$ (S = 1.2)
$\phi_\epsilon = (2\phi_{+-} + \phi_{00}) / 3$	$(43.5 \pm 0.5)^\circ$ (S = 1.3)
CP asymmetry A in $K_L^0 \rightarrow \pi^+ \pi^- e^+ e^-$	$(13.7 \pm 1.5)\%$
$\beta_{CP}$ from $K_L^0 \rightarrow e^+ e^- e^+ e^-$	$-0.19 \pm 0.07$
$\gamma_{CP}$ from $K_L^0 \rightarrow e^+ e^- e^+ e^-$	$0.01 \pm 0.11$ (S = 1.6)
parameters for $K_L^0 \rightarrow \pi^+ \pi^- \gamma$ decay	
$ \eta_{+-\gamma}  =  A(K_L^0 \rightarrow \pi^+ \pi^- \gamma, CP \text{ violating})  / A(K_S^0 \rightarrow \pi^+ \pi^- \gamma) $	$(2.35 \pm 0.07) \times 10^{-3}$
$\phi_{+-\gamma}$ = phase of $\eta_{+-\gamma}$	$(44 \pm 4)^\circ$
$\Gamma(K_L^0 \rightarrow \pi^+ \pi^-) / \Gamma_{\text{total}}$	$[i] \ (1.967 \pm 0.010) \times 10^{-3}$ (S = 1.5)
$\Gamma(K_L^0 \rightarrow \pi^0 \pi^0) / \Gamma_{\text{total}}$	$(8.64 \pm 0.06) \times 10^{-4}$ (S = 1.8)
$ACP(B^+ \rightarrow D_{CP(+1)} K^+)$	$0.170 \pm 0.033$ (S = 1.2)
$A_{ADS}(B^+ \rightarrow D K^+)$	$-0.52 \pm 0.15$
$ACP(B^+ \rightarrow \eta K^+)$	$-0.37 \pm 0.08$
$ACP(B^+ \rightarrow f_2(1270) K^+)$	$-0.68^{+0.19}_{-0.17}$
$ACP(B^+ \rightarrow \rho^0 K^+)$	$0.37 \pm 0.10$
$ACP(B^+ \rightarrow f_0(1370) \pi^+)$	$0.72 \pm 0.22$
$\gamma(B^+ \rightarrow D^{(*)0} K^{(*)+})$	$(70 \pm 9)^\circ$

## CPT INVARIANCE

$(m_{W^+} - m_{W^-}) / m_{\text{average}}$	$-0.002 \pm 0.007$
$(m_{e^+} - m_{e^-}) / m_{\text{average}}$	$< 8 \times 10^{-9}$ , CL = 90%
$ q_{e^+} + q_{e^-}  / e$	$< 4 \times 10^{-8}$
$(g_{e^+} - g_{e^-}) / g_{\text{average}}$	$(-0.5 \pm 2.1) \times 10^{-12}$
$(\tau_{\mu^+} - \tau_{\mu^-}) / \tau_{\text{average}}$	$(2 \pm 8) \times 10^{-5}$
$(g_{\mu^+} - g_{\mu^-}) / g_{\text{average}}$	$(-0.11 \pm 0.12) \times 10^{-8}$
$(m_{\tau^+} - m_{\tau^-}) / m_{\text{average}}$	$< 2.8 \times 10^{-4}$ , CL = 90%
$(m_t - m_{\overline{t}}) / m_{\text{average}}$	$-0.2 \pm 0.5$ GeV (S = 1.1)
$(m_{\pi^+} - m_{\pi^-}) / m_{\text{average}}$	$(2 \pm 5) \times 10^{-4}$
$(\tau_{\pi^+} - \tau_{\pi^-}) / \tau_{\text{average}}$	$(6 \pm 7) \times 10^{-4}$
$(m_{K^+} - m_{K^-}) / m_{\text{average}}$	$(-0.6 \pm 1.8) \times 10^{-4}$
$(\tau_{K^+} - \tau_{K^-}) / \tau_{\text{average}}$	$(0.10 \pm 0.09)\%$ (S = 1.2)
$K^\pm \rightarrow \mu^\pm \nu_\mu$ rate difference/sum	$(-0.27 \pm 0.21)\%$
$K^\pm \rightarrow \pi^\pm \pi^0$ rate difference/sum	$[j] \ (0.4 \pm 0.6)\%$
$\delta$ in $K^0 - \overline{K}^0$ mixing	
real part of $\delta$	$(2.5 \pm 2.3) \times 10^{-4}$
imaginary part of $\delta$	$(-1.5 \pm 1.6) \times 10^{-5}$
$\text{Re}(y)$ , $K_{e3}$ parameter	$(0.4 \pm 2.5) \times 10^{-3}$
$\text{Re}(x_-)$ , $K_{e3}$ parameter	$(-2.9 \pm 2.0) \times 10^{-3}$
$ m_{K^0} - m_{\overline{K}^0}  / m_{\text{average}}$	$[k] \ < 6 \times 10^{-19}$ , CL = 90%
$(\Gamma_{K^0} - \Gamma_{\overline{K}^0}) / m_{\text{average}}$	$(8 \pm 8) \times 10^{-18}$
phase difference $\phi_{00} - \phi_{+-}$	$(0.34 \pm 0.32)^\circ$
$\text{Re}(\frac{2}{3}\eta_{+-} + \frac{1}{3}\eta_{00}) - \frac{A_L}{2}$	$(-3 \pm 35) \times 10^{-6}$
$ACPT(D^0 \rightarrow K^- \pi^+)$	$0.008 \pm 0.008$
$\Delta S_{CPT}^+(S_{\ell^+, K_S^0}^- - S_{\ell^+, K_S^0}^+)$	$0.16 \pm 0.23$
$\Delta S_{CPT}^-(S_{\ell^+, K_S^0}^+ - S_{\ell^+, K_S^0}^-)$	$-0.03 \pm 0.14$
$\Delta C_{CPT}^+(C_{\ell^+, K_S^0}^- - C_{\ell^+, K_S^0}^+)$	$0.14 \pm 0.17$
$\Delta C_{CPT}^-(C_{\ell^+, K_S^0}^+ - C_{\ell^+, K_S^0}^-)$	$0.03 \pm 0.14$
$ m_{\overline{p}} - m_p  / m_p$	$[l] \ < 7 \times 10^{-10}$ , CL = 90%
$(\frac{q_p}{m_p} - \frac{q_{\overline{p}}}{m_{\overline{p}}}) / \frac{q_p}{m_p}$	$(-9 \pm 9) \times 10^{-11}$
$ q_p + q_{\overline{p}}  / e$	$[j] \ < 7 \times 10^{-10}$ , CL = 90%
$(\mu_p + \mu_{\overline{p}}) / \mu_p$	$(0 \pm 5) \times 10^{-6}$
$(m_n - m_{\overline{n}}) / m_n$	$(9 \pm 6) \times 10^{-5}$
$(m_\Lambda - m_{\overline{\Lambda}}) / m_\Lambda$	$(-0.1 \pm 1.1) \times 10^{-5}$ (S = 1.6)
$(\tau_\Lambda - \tau_{\overline{\Lambda}}) / \tau_\Lambda$	$-0.001 \pm 0.009$
$(\tau_{\Sigma^+} - \tau_{\overline{\Sigma}^-}) / \tau_{\Sigma^+}$	$-0.0006 \pm 0.0012$

Unless otherwise stated, limits are given at the 90% confidence level, while errors are given as  $\pm 1$  standard deviation.

# Tests of Conservation Laws

$(\mu_{\Sigma^+} + \mu_{\Sigma^-}) / \mu_{\Sigma^+}$	$0.014 \pm 0.015$
$(m_{\Xi^-} - m_{\Xi^+}) / m_{\Xi^-}$	$(-3 \pm 9) \times 10^{-5}$
$(\tau_{\Xi^-} - \tau_{\Xi^+}) / \tau_{\Xi^-}$	$-0.01 \pm 0.07$
$(\mu_{\Xi^-} + \mu_{\Xi^+}) /  \mu_{\Xi^-} $	$+0.01 \pm 0.05$
$(m_{\Omega^-} - m_{\Omega^+}) / m_{\Omega^-}$	$(-1 \pm 8) \times 10^{-5}$
$(\tau_{\Omega^-} - \tau_{\Omega^+}) / \tau_{\Omega^-}$	$0.00 \pm 0.05$

## TESTS OF NUMBER CONSERVATION LAWS

### LEPTON FAMILY NUMBER

Lepton family number conservation means separate conservation of each of  $L_e$ ,  $L_\mu$ ,  $L_\tau$ .

$\Gamma(Z \rightarrow e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 7.5 \times 10^{-7}$ , CL = 95%
$\Gamma(Z \rightarrow e^\pm \tau^\mp) / \Gamma_{\text{total}}$	$[n] < 9.8 \times 10^{-6}$ , CL = 95%
$\Gamma(Z \rightarrow \mu^\pm \tau^\mp) / \Gamma_{\text{total}}$	$[n] < 1.2 \times 10^{-5}$ , CL = 95%
$\sigma(e^+ e^- \rightarrow e^\pm \tau^\mp) / \sigma(e^+ e^- \rightarrow \mu^\pm \mu^-)$	$< 8.9 \times 10^{-6}$ , CL = 95%
$\sigma(e^+ e^- \rightarrow \mu^\pm \tau^\mp) / \sigma(e^+ e^- \rightarrow \mu^\pm \mu^-)$	$< 4.0 \times 10^{-6}$ , CL = 95%
limit on $\mu^- \rightarrow e^-$ conversion $\sigma(\mu^- 32\text{S} \rightarrow e^- 32\text{S}) / \sigma(\mu^- 32\text{S} \rightarrow \nu_\mu 32\text{P}^*)$	$< 7 \times 10^{-11}$ , CL = 90%
$\sigma(\mu^- \text{Ti} \rightarrow e^- \text{Ti}) / \sigma(\mu^- \text{Ti} \rightarrow \text{capture})$	$< 4.3 \times 10^{-12}$ , CL = 90%
$\sigma(\mu^- \text{Pb} \rightarrow e^- \text{Pb}) / \sigma(\mu^- \text{Pb} \rightarrow \text{capture})$	$< 4.6 \times 10^{-11}$ , CL = 90%
limit on muonium $\rightarrow$ antimuonium conversion $R_g = G_C / G_F$	$< 0.0030$ , CL = 90%
$\Gamma(\mu^- \rightarrow e^- \nu_e \bar{\nu}_\mu) / \Gamma_{\text{total}}$	$[a] < 1.2 \times 10^{-2}$ , CL = 90%
$\Gamma(\mu^- \rightarrow e^- \gamma) / \Gamma_{\text{total}}$	$< 5.7 \times 10^{-13}$ , CL = 90%
$\Gamma(\mu^- \rightarrow e^- e^+ e^-) / \Gamma_{\text{total}}$	$< 1.0 \times 10^{-12}$ , CL = 90%
$\Gamma(\mu^- \rightarrow e^- 2\gamma) / \Gamma_{\text{total}}$	$< 7.2 \times 10^{-11}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \gamma) / \Gamma_{\text{total}}$	$< 3.3 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \gamma) / \Gamma_{\text{total}}$	$< 4.4 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \pi^0) / \Gamma_{\text{total}}$	$< 8.0 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \pi^0) / \Gamma_{\text{total}}$	$< 1.1 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- K_S^0) / \Gamma_{\text{total}}$	$< 2.6 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- K_S^0) / \Gamma_{\text{total}}$	$< 2.3 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \eta) / \Gamma_{\text{total}}$	$< 9.2 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \eta) / \Gamma_{\text{total}}$	$< 6.5 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \rho^0) / \Gamma_{\text{total}}$	$< 1.8 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \rho^0) / \Gamma_{\text{total}}$	$< 1.2 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \omega) / \Gamma_{\text{total}}$	$< 4.8 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \omega) / \Gamma_{\text{total}}$	$< 4.7 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- K^*(892)^0) / \Gamma_{\text{total}}$	$< 3.2 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- K^*(892)^0) / \Gamma_{\text{total}}$	$< 5.9 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \bar{K}^*(892)^0) / \Gamma_{\text{total}}$	$< 3.4 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \bar{K}^*(892)^0) / \Gamma_{\text{total}}$	$< 7.0 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \eta'(958)) / \Gamma_{\text{total}}$	$< 1.6 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \eta'(958)) / \Gamma_{\text{total}}$	$< 1.3 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- f_0(980) \rightarrow e^- \pi^+ \pi^-) / \Gamma_{\text{total}}$	$< 3.2 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-) / \Gamma_{\text{total}}$	$< 3.4 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \phi) / \Gamma_{\text{total}}$	$< 3.1 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \phi) / \Gamma_{\text{total}}$	$< 8.4 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- e^+ e^-) / \Gamma_{\text{total}}$	$< 2.7 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \mu^+ \mu^-) / \Gamma_{\text{total}}$	$< 2.7 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^+ \mu^- \mu^-) / \Gamma_{\text{total}}$	$< 1.7 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^+ e^- e^-) / \Gamma_{\text{total}}$	$< 1.8 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^+ \mu^+ e^-) / \Gamma_{\text{total}}$	$< 1.5 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^+ \mu^+ \mu^-) / \Gamma_{\text{total}}$	$< 2.1 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \pi^+ \pi^-) / \Gamma_{\text{total}}$	$< 2.3 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \pi^+ \pi^-) / \Gamma_{\text{total}}$	$< 2.1 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \pi^+ K^-) / \Gamma_{\text{total}}$	$< 3.7 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \pi^- K^+) / \Gamma_{\text{total}}$	$< 3.1 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- K_S^0 K_S^0) / \Gamma_{\text{total}}$	$< 7.1 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- K^+ K^-) / \Gamma_{\text{total}}$	$< 3.4 \times 10^{-8}$ , CL = 90%

$\Gamma(\tau^- \rightarrow \mu^- \pi^+ K^-) / \Gamma_{\text{total}}$	$< 8.6 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \pi^- K^+) / \Gamma_{\text{total}}$	$< 4.5 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- K_S^0 K_S^0) / \Gamma_{\text{total}}$	$< 8.0 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- K^+ K^-) / \Gamma_{\text{total}}$	$< 4.4 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \pi^0 \pi^0) / \Gamma_{\text{total}}$	$< 6.5 \times 10^{-6}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \pi^0 \pi^0) / \Gamma_{\text{total}}$	$< 1.4 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \eta \eta) / \Gamma_{\text{total}}$	$< 3.5 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \eta \eta) / \Gamma_{\text{total}}$	$< 6.0 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \pi^0 \eta) / \Gamma_{\text{total}}$	$< 2.4 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^- \pi^0 \eta) / \Gamma_{\text{total}}$	$< 2.2 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow e^- \text{light boson}) / \Gamma_{\text{total}}$	$< 2.7 \times 10^{-3}$ , CL = 95%
$\Gamma(\tau^- \rightarrow \mu^- \text{light boson}) / \Gamma_{\text{total}}$	$< 5 \times 10^{-3}$ , CL = 95%

### LEPTON FAMILY NUMBER VIOLATION IN NEUTRINOS

$\sin^2(\theta_{12})$	$0.304 \pm 0.014$
$\Delta m_{21}^2$	$(7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2$
$\sin^2(\theta_{23})$ (normal mass hierarchy)	$0.51 \pm 0.05$
$\sin^2(\theta_{23})$ (inverted mass hierarchy)	$0.50 \pm 0.05$
$\Delta m_{32}^2$ (normal mass hierarchy)	$[p] (2.44 \pm 0.06) \times 10^{-3} \text{ eV}^2$
$\Delta m_{32}^2$ (inverted mass hierarchy)	$[p] (2.51 \pm 0.06) \times 10^{-3} \text{ eV}^2$
$\sin^2(\theta_{13})$	$(2.19 \pm 0.12) \times 10^{-2}$
$\Gamma(\pi^+ \rightarrow \mu^+ \nu_e) / \Gamma_{\text{total}}$	$[q] < 8.0 \times 10^{-3}$ , CL = 90%
$\Gamma(\pi^+ \rightarrow \mu^- e^+ e^- \nu) / \Gamma_{\text{total}}$	$< 1.6 \times 10^{-6}$ , CL = 90%
$\Gamma(\pi^0 \rightarrow \mu^+ e^-) / \Gamma_{\text{total}}$	$< 3.8 \times 10^{-10}$ , CL = 90%
$\Gamma(\pi^0 \rightarrow \mu^- e^+) / \Gamma_{\text{total}}$	$< 3.4 \times 10^{-9}$ , CL = 90%
$\Gamma(\pi^0 \rightarrow \mu^+ e^- + \mu^- e^+) / \Gamma_{\text{total}}$	$< 3.6 \times 10^{-10}$ , CL = 90%
$\Gamma(\eta \rightarrow \mu^+ e^- + \mu^- e^+) / \Gamma_{\text{total}}$	$< 6 \times 10^{-6}$ , CL = 90%
$\Gamma(\eta'(958) \rightarrow e \mu) / \Gamma_{\text{total}}$	$< 4.7 \times 10^{-4}$ , CL = 90%
$\Gamma(\phi(1020) \rightarrow e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$< 2 \times 10^{-6}$ , CL = 90%
$\Gamma(K^+ \rightarrow \mu^- \nu e^+ e^+) / \Gamma_{\text{total}}$	$< 2.1 \times 10^{-8}$ , CL = 90%
$\Gamma(K^+ \rightarrow \mu^+ \nu_e) / \Gamma_{\text{total}}$	$[q] < 4 \times 10^{-3}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^+ \mu^+ e^-) / \Gamma_{\text{total}}$	$< 1.3 \times 10^{-11}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^+ \mu^- e^+) / \Gamma_{\text{total}}$	$< 5.2 \times 10^{-10}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 4.7 \times 10^{-12}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow e^\pm e^\pm \mu^\mp \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 4.12 \times 10^{-11}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \mu^\pm e^\mp) / \Gamma_{\text{total}}$	$[n] < 7.6 \times 10^{-11}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \pi^0 \mu^\pm e^\mp) / \Gamma_{\text{total}}$	$< 1.7 \times 10^{-10}$ , CL = 90%
$\Gamma(D^+ \rightarrow \pi^+ e^+ \mu^-) / \Gamma_{\text{total}}$	$< 2.9 \times 10^{-6}$ , CL = 90%
$\Gamma(D^+ \rightarrow \pi^+ e^- \mu^+) / \Gamma_{\text{total}}$	$< 3.6 \times 10^{-6}$ , CL = 90%
$\Gamma(D^+ \rightarrow K^+ e^+ \mu^-) / \Gamma_{\text{total}}$	$< 1.2 \times 10^{-6}$ , CL = 90%
$\Gamma(D^+ \rightarrow K^+ e^- \mu^+) / \Gamma_{\text{total}}$	$< 2.8 \times 10^{-6}$ , CL = 90%
$\Gamma(D^0 \rightarrow \mu^\pm e^\mp) / \Gamma_{\text{total}}$	$[n] < 2.6 \times 10^{-7}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^0 e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 8.6 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \eta e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 1.0 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^+ \pi^- e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 1.5 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \rho^0 e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 4.9 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \omega e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 1.2 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow K^- K^+ e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 1.8 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \phi e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 3.4 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \bar{K}^0 e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 1.0 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow K^- \pi^+ e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 5.53 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \bar{K}^*(892)^0 e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$[n] < 8.3 \times 10^{-5}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow \pi^+ e^+ \mu^-) / \Gamma_{\text{total}}$	$< 1.2 \times 10^{-5}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow \pi^+ e^- \mu^+) / \Gamma_{\text{total}}$	$< 2.0 \times 10^{-5}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow K^+ e^+ \mu^-) / \Gamma_{\text{total}}$	$< 1.4 \times 10^{-5}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow K^+ e^- \mu^+) / \Gamma_{\text{total}}$	$< 9.7 \times 10^{-6}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^+ \mu^-) / \Gamma_{\text{total}}$	$< 6.4 \times 10^{-3}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^- \mu^+) / \Gamma_{\text{total}}$	$< 6.4 \times 10^{-3}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$< 1.7 \times 10^{-7}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^+ \tau^-) / \Gamma_{\text{total}}$	$< 7.4 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^- \tau^+) / \Gamma_{\text{total}}$	$< 2.0 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^\pm \tau^\mp) / \Gamma_{\text{total}}$	$< 7.5 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ \mu^+ \tau^-) / \Gamma_{\text{total}}$	$< 6.2 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ \mu^- \tau^+) / \Gamma_{\text{total}}$	$< 4.5 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ \mu^\pm \tau^\mp) / \Gamma_{\text{total}}$	$< 7.2 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ e^+ \mu^-) / \Gamma_{\text{total}}$	$< 9.1 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ e^- \mu^+) / \Gamma_{\text{total}}$	$< 1.3 \times 10^{-7}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ e^\pm \mu^\mp) / \Gamma_{\text{total}}$	$< 9.1 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ e^+ \tau^-) / \Gamma_{\text{total}}$	$< 4.3 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ e^- \tau^+) / \Gamma_{\text{total}}$	$< 1.5 \times 10^{-5}$ , CL = 90%

Unless otherwise stated, limits are given at the 90% confidence level, while errors are given as  $\pm 1$  standard deviation.



# Tests of Conservation Laws

$\Gamma(B^+ \rightarrow K^+ e^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<3.0 \times 10^{-5}$ , CL = 90%	$\Gamma(D^+ \rightarrow \rho^- 2\mu^+)/\Gamma_{\text{total}}$	$<5.6 \times 10^{-4}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<4.5 \times 10^{-5}$ , CL = 90%	$\Gamma(D^+ \rightarrow K^- 2e^+)/\Gamma_{\text{total}}$	$<9 \times 10^{-7}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<2.8 \times 10^{-5}$ , CL = 90%	$\Gamma(D^+ \rightarrow K^- 2\mu^+)/\Gamma_{\text{total}}$	$<1.0 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<4.8 \times 10^{-5}$ , CL = 90%	$\Gamma(D^+ \rightarrow K^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<1.9 \times 10^{-6}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^*(892)^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-6}$ , CL = 90%	$\Gamma(D^+ \rightarrow K^*(892)^- 2\mu^+)/\Gamma_{\text{total}}$	$<8.5 \times 10^{-4}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^*(892)^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<9.9 \times 10^{-7}$ , CL = 90%	$\Gamma(D^0 \rightarrow 2\pi^- 2e^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<1.12 \times 10^{-4}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^*(892)^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-6}$ , CL = 90%	$\Gamma(D^0 \rightarrow 2\pi^- 2\mu^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<2.9 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow e^\pm \mu^\mp)/\Gamma_{\text{total}}$	[n] $<2.8 \times 10^{-9}$ , CL = 90%	$\Gamma(D^0 \rightarrow K^- \pi^- 2e^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<2.06 \times 10^{-4}$ , CL = 90%
$\Gamma(B^0 \rightarrow \pi^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-7}$ , CL = 90%	$\Gamma(D^0 \rightarrow K^- \pi^- 2\mu^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<3.9 \times 10^{-4}$ , CL = 90%
$\Gamma(B^0 \rightarrow K^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-7}$ , CL = 90%	$\Gamma(D^0 \rightarrow 2K^- 2e^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<1.52 \times 10^{-4}$ , CL = 90%
$\Gamma(B^0 \rightarrow K^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<5.3 \times 10^{-7}$ , CL = 90%	$\Gamma(D^0 \rightarrow 2K^- 2\mu^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<9.4 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow K^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<3.4 \times 10^{-7}$ , CL = 90%	$\Gamma(D^0 \rightarrow \pi^- \pi^- e^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<7.9 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow K^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<5.8 \times 10^{-7}$ , CL = 90%	$\Gamma(D^0 \rightarrow K^- \pi^- e^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<2.18 \times 10^{-4}$ , CL = 90%
$\Gamma(B^0 \rightarrow e^\pm \tau^\mp)/\Gamma_{\text{total}}$	[n] $<2.8 \times 10^{-5}$ , CL = 90%	$\Gamma(D^0 \rightarrow 2K^- e^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$	$<5.7 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	[n] $<2.2 \times 10^{-5}$ , CL = 90%	$\Gamma(D^0 \rightarrow p e^-)/\Gamma_{\text{total}}$	[r] $<1.0 \times 10^{-5}$ , CL = 90%
$\Gamma(B \rightarrow s e^\pm \mu^\mp)/\Gamma_{\text{total}}$	[n] $<2.2 \times 10^{-5}$ , CL = 90%	$\Gamma(D^0 \rightarrow \bar{p} e^+)/\Gamma_{\text{total}}$	[s] $<1.1 \times 10^{-5}$ , CL = 90%
$\Gamma(B \rightarrow \pi e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<9.2 \times 10^{-8}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow \pi^- 2e^+)/\Gamma_{\text{total}}$	$<4.1 \times 10^{-6}$ , CL = 90%
$\Gamma(B \rightarrow \rho e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-6}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow \pi^- 2\mu^+)/\Gamma_{\text{total}}$	$<1.2 \times 10^{-7}$ , CL = 90%
$\Gamma(B \rightarrow K e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<3.8 \times 10^{-8}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow \pi^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<8.4 \times 10^{-6}$ , CL = 90%
$\Gamma(B \rightarrow K^*(892) e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<5.1 \times 10^{-7}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow K^- 2e^+)/\Gamma_{\text{total}}$	$<5.2 \times 10^{-6}$ , CL = 90%
$\Gamma(B_S^0 \rightarrow e^\pm \mu^\mp)/\Gamma_{\text{total}}$	[n] $<1.1 \times 10^{-8}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow K^- 2\mu^+)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-5}$ , CL = 90%
$\Gamma(J/\psi(1S) \rightarrow e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<1.6 \times 10^{-7}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow K^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<6.1 \times 10^{-6}$ , CL = 90%
$\Gamma(J/\psi(1S) \rightarrow e^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<8.3 \times 10^{-6}$ , CL = 90%	$\Gamma(D_S^+ \rightarrow K^*(892)^- 2\mu^+)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-3}$ , CL = 90%
$\Gamma(J/\psi(1S) \rightarrow \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<2.0 \times 10^{-6}$ , CL = 90%	$\Gamma(B^+ \rightarrow \pi^- e^+ e^+)/\Gamma_{\text{total}}$	$<2.3 \times 10^{-8}$ , CL = 90%
$\Gamma(\Upsilon(1S) \rightarrow \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<6.0 \times 10^{-6}$ , CL = 95%	$\Gamma(B^+ \rightarrow \pi^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<4.0 \times 10^{-9}$ , CL = 95%
$\Gamma(\Upsilon(2S) \rightarrow e^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-6}$ , CL = 90%	$\Gamma(B^+ \rightarrow \pi^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<1.5 \times 10^{-7}$ , CL = 90%
$\Gamma(\Upsilon(2S) \rightarrow \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-6}$ , CL = 90%	$\Gamma(B^+ \rightarrow \rho^- e^+ e^+)/\Gamma_{\text{total}}$	$<1.7 \times 10^{-7}$ , CL = 90%
$\Gamma(\Upsilon(3S) \rightarrow e^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<4.2 \times 10^{-6}$ , CL = 90%	$\Gamma(B^+ \rightarrow \rho^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<4.2 \times 10^{-7}$ , CL = 90%
$\Gamma(\Upsilon(3S) \rightarrow \mu^\pm \tau^\mp)/\Gamma_{\text{total}}$	$<3.1 \times 10^{-6}$ , CL = 90%	$\Gamma(B^+ \rightarrow \rho^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<4.7 \times 10^{-7}$ , CL = 90%
$\Gamma(\Lambda_C^+ \rightarrow p e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<9.9 \times 10^{-6}$ , CL = 90%	$\Gamma(B^+ \rightarrow K^- e^+ e^+)/\Gamma_{\text{total}}$	$<3.0 \times 10^{-8}$ , CL = 90%
$\Gamma(\Lambda_C^+ \rightarrow p e^\pm \mu^\mp)/\Gamma_{\text{total}}$	$<1.9 \times 10^{-5}$ , CL = 90%	$\Gamma(B^+ \rightarrow K^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<4.1 \times 10^{-8}$ , CL = 90%
<b>TOTAL LEPTON NUMBER</b>			
Violation of total lepton number conservation also implies violation of lepton family number conservation.			
$\Gamma(Z \rightarrow p e)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$ , CL = 95%	$\Gamma(B^+ \rightarrow K^*(892)^- e^+ e^+)/\Gamma_{\text{total}}$	$<4.0 \times 10^{-7}$ , CL = 90%
$\Gamma(Z \rightarrow p \mu)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$ , CL = 95%	$\Gamma(B^+ \rightarrow K^*(892)^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<5.9 \times 10^{-7}$ , CL = 90%
limit on $\mu^- \rightarrow e^+$ conversion		$\Gamma(B^+ \rightarrow K^*(892)^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<3.0 \times 10^{-7}$ , CL = 90%
$\sigma(\mu^- 32S \rightarrow e^+ 32Si^*) / \sigma(\mu^- 32S \rightarrow \nu_\mu 32P^*)$	$<9 \times 10^{-10}$ , CL = 90%	$\Gamma(B^+ \rightarrow D^- e^+ e^+)/\Gamma_{\text{total}}$	$<2.6 \times 10^{-6}$ , CL = 90%
$\sigma(\mu^- 127I \rightarrow e^+ 127Sb^*) / \sigma(\mu^- 127I \rightarrow \text{anything})$	$<3 \times 10^{-10}$ , CL = 90%	$\Gamma(B^+ \rightarrow D^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$ , CL = 90%
$\sigma(\mu^- Ti \rightarrow e^+ Ca) / \sigma(\mu^- Ti \rightarrow \text{capture})$	$<3.6 \times 10^{-11}$ , CL = 90%	$\Gamma(B^+ \rightarrow D^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<6.9 \times 10^{-7}$ , CL = 95%
$\Gamma(\tau^- \rightarrow e^+ \pi^- \pi^-)/\Gamma_{\text{total}}$	$<2.0 \times 10^{-8}$ , CL = 90%	$\Gamma(B^+ \rightarrow D^{*-} \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<2.4 \times 10^{-6}$ , CL = 95%
$\Gamma(\tau^- \rightarrow \mu^+ \pi^- \pi^-)/\Gamma_{\text{total}}$	$<3.9 \times 10^{-8}$ , CL = 90%	$\Gamma(B^+ \rightarrow D_S^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<5.8 \times 10^{-7}$ , CL = 95%
$\Gamma(\tau^- \rightarrow e^+ \pi^- K^-)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-8}$ , CL = 90%	$\Gamma(B^+ \rightarrow \bar{D}^0 \pi^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<1.5 \times 10^{-6}$ , CL = 95%
$\Gamma(\tau^- \rightarrow e^+ K^- K^-)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-8}$ , CL = 90%	$\Gamma(B^+ \rightarrow \Lambda^0 \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^+ \pi^- K^-)/\Gamma_{\text{total}}$	$<4.8 \times 10^{-8}$ , CL = 90%	$\Gamma(B^+ \rightarrow \Lambda^0 e^+)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \mu^+ K^- K^-)/\Gamma_{\text{total}}$	$<4.7 \times 10^{-8}$ , CL = 90%	$\Gamma(B^+ \rightarrow \bar{\Lambda}^0 \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow p \mu^- \mu^-)/\Gamma_{\text{total}}$	$<4.4 \times 10^{-7}$ , CL = 90%	$\Gamma(B^+ \rightarrow \bar{\Lambda}^0 e^+)/\Gamma_{\text{total}}$	$<8 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-7}$ , CL = 90%	$\Gamma(B^0 \rightarrow \Lambda_C^+ \mu^-)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-6}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \gamma)/\Gamma_{\text{total}}$	$<3.5 \times 10^{-6}$ , CL = 90%	$\Gamma(B^0 \rightarrow \Lambda_C^+ e^-)/\Gamma_{\text{total}}$	$<4 \times 10^{-6}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \pi^0)/\Gamma_{\text{total}}$	$<1.5 \times 10^{-5}$ , CL = 90%	$\Gamma(\Lambda \rightarrow \pi^+ e^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} 2\pi^0)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-5}$ , CL = 90%	$\Gamma(\Lambda \rightarrow \pi^+ \mu^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \eta)/\Gamma_{\text{total}}$	$<8.9 \times 10^{-6}$ , CL = 90%	$\Gamma(\Lambda \rightarrow \pi^- e^+)/\Gamma_{\text{total}}$	$<4 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \pi^0 \eta)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-5}$ , CL = 90%	$\Gamma(\Lambda \rightarrow \pi^- \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \Lambda \pi^-)/\Gamma_{\text{total}}$	$<7.2 \times 10^{-8}$ , CL = 90%	$\Gamma(\Lambda \rightarrow K^+ e^-)/\Gamma_{\text{total}}$	$<2 \times 10^{-6}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{\Lambda} \pi^-)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-7}$ , CL = 90%	$\Gamma(\Lambda \rightarrow K^+ \mu^-)/\Gamma_{\text{total}}$	$<3 \times 10^{-6}$ , CL = 90%
$t_{1/2}(^{76}\text{Ge} \rightarrow ^{76}\text{Se} + 2 e^-)$	$>1.9 \times 10^{25}$ yr, CL = 90%	$\Gamma(\Lambda \rightarrow K^- e^+)/\Gamma_{\text{total}}$	$<2 \times 10^{-6}$ , CL = 90%
$\Gamma(\pi^+ \rightarrow \mu^+ \bar{\nu}_e)/\Gamma_{\text{total}}$	[q] $<1.5 \times 10^{-3}$ , CL = 90%	$\Gamma(\Lambda \rightarrow K^- \mu^+)/\Gamma_{\text{total}}$	$<3 \times 10^{-6}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^- \mu^+ e^+)/\Gamma_{\text{total}}$	$<5.0 \times 10^{-10}$ , CL = 90%	$\Gamma(\Lambda \rightarrow \Sigma^0 \nu)/\Gamma_{\text{total}}$	$<2 \times 10^{-5}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^- e^+ e^+)/\Gamma_{\text{total}}$	$<6.4 \times 10^{-10}$ , CL = 90%	$\Gamma(\Xi^- \rightarrow p \mu^- \mu^-)/\Gamma_{\text{total}}$	$<4 \times 10^{-8}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	[q] $<1.1 \times 10^{-9}$ , CL = 90%	$\Gamma(\Lambda_C^+ \rightarrow \bar{p} 2e^+)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-6}$ , CL = 90%
$\Gamma(K^+ \rightarrow \mu^+ \bar{\nu}_e)/\Gamma_{\text{total}}$	[q] $<3.3 \times 10^{-3}$ , CL = 90%	$\Gamma(\Lambda_C^+ \rightarrow \bar{p} 2\mu^+)/\Gamma_{\text{total}}$	$<9.4 \times 10^{-6}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^0 e^+ \bar{\nu}_e)/\Gamma_{\text{total}}$	$<3 \times 10^{-3}$ , CL = 90%	$\Gamma(\Lambda_C^+ \rightarrow \bar{p} e^+ \mu^+)/\Gamma_{\text{total}}$	$<1.6 \times 10^{-5}$ , CL = 90%
$\Gamma(D^+ \rightarrow \pi^- 2e^+)/\Gamma_{\text{total}}$	$<1.1 \times 10^{-6}$ , CL = 90%	$\Gamma(\Lambda_C^+ \rightarrow \Sigma^- \mu^+ \mu^+)/\Gamma_{\text{total}}$	$<7.0 \times 10^{-4}$ , CL = 90%
$\Gamma(D^+ \rightarrow \pi^- 2\mu^+)/\Gamma_{\text{total}}$	$<2.2 \times 10^{-8}$ , CL = 90%		
$\Gamma(D^+ \rightarrow \pi^- e^+ \mu^+)/\Gamma_{\text{total}}$	$<2.0 \times 10^{-6}$ , CL = 90%		

Unless otherwise stated, limits are given at the 90% confidence level, while errors are given as  $\pm 1$  standard deviation.

# Tests of Conservation Laws

## BARYON NUMBER

$\Gamma(Z \rightarrow p e)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$ , CL = 95%
$\Gamma(Z \rightarrow p \mu)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$ , CL = 95%
$\Gamma(\tau^- \rightarrow p \mu^- \mu^-)/\Gamma_{\text{total}}$	$<4.4 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-7}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \gamma)/\Gamma_{\text{total}}$	$<3.5 \times 10^{-6}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \pi^0)/\Gamma_{\text{total}}$	$<1.5 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} 2\pi^0)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \eta)/\Gamma_{\text{total}}$	$<8.9 \times 10^{-6}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \pi^0 \eta)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-5}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \Lambda \pi^-)/\Gamma_{\text{total}}$	$<7.2 \times 10^{-8}$ , CL = 90%
$\Gamma(\tau^- \rightarrow \bar{\Lambda} \pi^-)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-7}$ , CL = 90%
$\Gamma(D^0 \rightarrow p e^-)/\Gamma_{\text{total}}$	[r] $<1.0 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \bar{p} e^+)/\Gamma_{\text{total}}$	[s] $<1.1 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \Lambda^0 \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow \Lambda^0 e^+)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow \bar{\Lambda}^0 \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow \bar{\Lambda}^0 e^+)/\Gamma_{\text{total}}$	$<8 \times 10^{-8}$ , CL = 90%
$\Gamma(B^0 \rightarrow \Lambda_c^+ \mu^-)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-6}$ , CL = 90%
$\Gamma(B^0 \rightarrow \Lambda_c^+ e^-)/\Gamma_{\text{total}}$	$<4 \times 10^{-6}$ , CL = 90%

$p$  mean life [t]  $>2.1 \times 10^{29}$  years, CL = 90%

A few examples of proton or bound neutron decay follow. For limits on many other nucleon decay channels, see the Baryon Summary Table.

$\tau(N \rightarrow e^+ \pi)$	$> 2000 (n), > 8200 (p) \times 10^{30}$ years, CL = 90%
$\tau(N \rightarrow \mu^+ \pi)$	$> 1000 (n), > 6600 (p) \times 10^{30}$ years, CL = 90%
$\tau(N \rightarrow e^+ K)$	$> 17 (n), > 1000 (p) \times 10^{30}$ years, CL = 90%
$\tau(N \rightarrow \mu^+ K)$	$> 26 (n), > 1600 (p) \times 10^{30}$ years, CL = 90%
limit on $n\bar{n}$ oscillations (free $n$ )	$>0.86 \times 10^8$ s, CL = 90%
limit on $n\bar{n}$ oscillations (bound $n$ )	[u] $>1.3 \times 10^8$ s, CL = 90%
$\Gamma(\Lambda \rightarrow \pi^+ e^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-7}$ , CL = 90%
$\Gamma(\Lambda \rightarrow \pi^+ \mu^-)/\Gamma_{\text{total}}$	$<6 \times 10^{-7}$ , CL = 90%
$\Gamma(\Lambda \rightarrow \pi^- e^+)/\Gamma_{\text{total}}$	$<4 \times 10^{-7}$ , CL = 90%
$\Gamma(\Lambda \rightarrow \pi^- \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-7}$ , CL = 90%
$\Gamma(\Lambda \rightarrow K^+ e^-)/\Gamma_{\text{total}}$	$<2 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda \rightarrow K^+ \mu^-)/\Gamma_{\text{total}}$	$<3 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda \rightarrow K^- e^+)/\Gamma_{\text{total}}$	$<2 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda \rightarrow K^- \mu^+)/\Gamma_{\text{total}}$	$<3 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda \rightarrow K_S^0 \nu)/\Gamma_{\text{total}}$	$<2 \times 10^{-5}$ , CL = 90%
$\Gamma(\Lambda \rightarrow \bar{p} \pi^+)/\Gamma_{\text{total}}$	$<9 \times 10^{-7}$ , CL = 90%
$\Gamma(\Lambda_c^+ \rightarrow \bar{p} 2e^+)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda_c^+ \rightarrow \bar{p} 2\mu^+)/\Gamma_{\text{total}}$	$<9.4 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda_c^+ \rightarrow \bar{p} e^+ \mu^+)/\Gamma_{\text{total}}$	$<1.6 \times 10^{-5}$ , CL = 90%

## ELECTRIC CHARGE (Q)

$e \rightarrow \nu_e \gamma$ and astrophysical limits	[v] $>6.6 \times 10^{28}$ yr, CL = 90%
$\Gamma(n \rightarrow p \nu_e \bar{\nu}_e)/\Gamma_{\text{total}}$	$<8 \times 10^{-27}$ , CL = 68%

## $\Delta S = \Delta Q$ RULE

Violations allowed in second-order weak interactions.

$\Gamma(K^+ \rightarrow \pi^+ \pi^+ e^- \bar{\nu}_e)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-8}$ , CL = 90%
$\Gamma(K^+ \rightarrow \pi^+ \pi^+ \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$	$<3.0 \times 10^{-6}$ , CL = 95%
Re( $x_+$ ), $K_{e3}$ parameter	$(-0.9 \pm 3.0) \times 10^{-3}$
$x = A(\bar{K}^0 \rightarrow \pi^- \ell^+ \nu)/A(K^0 \rightarrow \pi^- \ell^+ \nu) = A(\Delta S = -\Delta Q)/A(\Delta S = \Delta Q)$	
real part of $x$	$-0.002 \pm 0.006$
imaginary part of $x$	$0.0012 \pm 0.0021$
$\Gamma(\Sigma^+ \rightarrow n \ell^+ \nu)/\Gamma(\Sigma^- \rightarrow n \ell^- \bar{\nu})$	$<0.043$
$\Gamma(\Sigma^+ \rightarrow n e^+ \nu_e)/\Gamma_{\text{total}}$	$<5 \times 10^{-6}$ , CL = 90%
$\Gamma(\Sigma^+ \rightarrow n \mu^+ \nu_\mu)/\Gamma_{\text{total}}$	$<3.0 \times 10^{-5}$ , CL = 90%
$\Gamma(\Xi^0 \rightarrow \Sigma^- e^+ \nu_e)/\Gamma_{\text{total}}$	$<9 \times 10^{-4}$ , CL = 90%
$\Gamma(\Xi^0 \rightarrow \Sigma^- \mu^+ \nu_\mu)/\Gamma_{\text{total}}$	$<9 \times 10^{-4}$ , CL = 90%

## $\Delta S = 2$ FORBIDDEN

Allowed in second-order weak interactions.

$\Gamma(\Xi^0 \rightarrow p \pi^-)/\Gamma_{\text{total}}$	$<8 \times 10^{-6}$ , CL = 90%
$\Gamma(\Xi^0 \rightarrow p e^- \bar{\nu}_e)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-3}$
$\Gamma(\Xi^0 \rightarrow p \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$	$<1.3 \times 10^{-3}$
$\Gamma(\Xi^- \rightarrow n \pi^-)/\Gamma_{\text{total}}$	$<1.9 \times 10^{-5}$ , CL = 90%
$\Gamma(\Xi^- \rightarrow n e^- \bar{\nu}_e)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-3}$ , CL = 90%
$\Gamma(\Xi^- \rightarrow n \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$	$<1.5 \times 10^{-2}$ , CL = 90%
$\Gamma(\Xi^- \rightarrow p \pi^- \pi^-)/\Gamma_{\text{total}}$	$<4 \times 10^{-4}$ , CL = 90%
$\Gamma(\Xi^- \rightarrow p \pi^- e^- \bar{\nu}_e)/\Gamma_{\text{total}}$	$<4 \times 10^{-4}$ , CL = 90%
$\Gamma(\Xi^- \rightarrow p \pi^- \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$	$<4 \times 10^{-4}$ , CL = 90%
$\Gamma(\Omega^- \rightarrow \Lambda \pi^-)/\Gamma_{\text{total}}$	$<2.9 \times 10^{-6}$ , CL = 90%

## $\Delta S = 2$ VIA MIXING

Allowed in second-order weak interactions, e.g. mixing.

$m_{K_L^0} - m_{K_S^0}$	$(0.5293 \pm 0.0009) \times 10^{10} \hbar \text{ s}^{-1}$ (S = 1.3)
$m_{K_L^0} - m_{K_S^0}$	$(3.484 \pm 0.006) \times 10^{-12} \text{ MeV}$

## $\Delta C = 2$ VIA MIXING

Allowed in second-order weak interactions, e.g. mixing.

$ m_{D_1^0} - m_{D_2^0}  = x\Gamma$	$(0.95_{-0.44}^{+0.41}) \times 10^{10} \hbar \text{ s}^{-1}$
$(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma = 2y$	$(1.29_{-0.18}^{+0.14}) \times 10^{-2}$

## $\Delta B = 2$ VIA MIXING

Allowed in second-order weak interactions, e.g. mixing.

$x_d$	$0.1875 \pm 0.0017$
$\Delta m_{B^0} = m_{B_H^0} - m_{B_L^0}$	$(0.5096 \pm 0.0034) \times 10^{12} \hbar \text{ s}^{-1}$
$x_d = \Delta m_{B^0}/\Gamma_{B^0}$	$0.775 \pm 0.006$
$\Delta m_{B_S^0} = m_{B_{SH}^0} - m_{B_{SL}^0}$	$(17.757 \pm 0.021) \times 10^{12} \hbar \text{ s}^{-1}$
$x_s = \Delta m_{B_S^0}/\Gamma_{B_S^0}$	$26.81 \pm 0.10$
$x_s$	$0.499308 \pm 0.000005$

## $\Delta S = 1$ WEAK NEUTRAL CURRENT FORBIDDEN

Allowed by higher-order electroweak interactions.

$\Gamma(K^+ \rightarrow \pi^+ e^+ e^-)/\Gamma_{\text{total}}$	$(3.00 \pm 0.09) \times 10^{-7}$
$\Gamma(K^+ \rightarrow \pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(9.4 \pm 0.6) \times 10^{-8}$ (S = 2.6)
$\Gamma(K^+ \rightarrow \pi^+ \nu \bar{\nu})/\Gamma_{\text{total}}$	$(1.7 \pm 1.1) \times 10^{-10}$
$\Gamma(K^+ \rightarrow \pi^+ \pi^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<4.3 \times 10^{-5}$ , CL = 90%
$\Gamma(K_S^0 \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<9 \times 10^{-9}$ , CL = 90%
$\Gamma(K_S^0 \rightarrow e^+ e^-)/\Gamma_{\text{total}}$	$<9 \times 10^{-9}$ , CL = 90%
$\Gamma(K_S^0 \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	[x] $(3.0_{-1.2}^{+1.5}) \times 10^{-9}$
$\Gamma(K_S^0 \rightarrow \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(2.9_{-1.2}^{+1.5}) \times 10^{-9}$
$\Gamma(K_L^0 \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(6.84 \pm 0.11) \times 10^{-9}$
$\Gamma(K_L^0 \rightarrow e^+ e^-)/\Gamma_{\text{total}}$	$(9_{-4}^{+6}) \times 10^{-12}$
$\Gamma(K_L^0 \rightarrow \pi^+ \pi^- e^+ e^-)/\Gamma_{\text{total}}$	[y] $(3.11 \pm 0.19) \times 10^{-7}$
$\Gamma(K_L^0 \rightarrow \pi^0 \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	$<6.6 \times 10^{-9}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<9.2 \times 10^{-11}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \mu^+ \mu^- e^+ e^-)/\Gamma_{\text{total}}$	$(2.69 \pm 0.27) \times 10^{-9}$
$\Gamma(K_L^0 \rightarrow e^+ e^- e^+ e^-)/\Gamma_{\text{total}}$	$(3.56 \pm 0.21) \times 10^{-8}$
$\Gamma(K_L^0 \rightarrow \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<3.8 \times 10^{-10}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	$<2.8 \times 10^{-10}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<2.6 \times 10^{-8}$ , CL = 90%
$\Gamma(K_L^0 \rightarrow \pi^0 \pi^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<8.1 \times 10^{-7}$ , CL = 90%
$\Gamma(\Sigma^+ \rightarrow p e^+ e^-)/\Gamma_{\text{total}}$	$<7 \times 10^{-6}$
$\Gamma(\Sigma^+ \rightarrow p \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(9_{-8}^{+9}) \times 10^{-8}$

# Tests of Conservation Laws

## $\Delta C = 1$ WEAK NEUTRAL CURRENT FORBIDDEN

Allowed by higher-order electroweak interactions.

$\Gamma(D^+ \rightarrow \pi^+ e^+ e^-)/\Gamma_{\text{total}}$	$<1.1 \times 10^{-6}$ , CL = 90%
$\Gamma(D^+ \rightarrow \pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<7.3 \times 10^{-8}$ , CL = 90%
$\Gamma(D^+ \rightarrow \rho^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<5.6 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$	$<2.2 \times 10^{-6}$ , CL = 90%
$\Gamma(D^0 \rightarrow e^+ e^-)/\Gamma_{\text{total}}$	$<7.9 \times 10^{-8}$ , CL = 90%
$\Gamma(D^0 \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<6.2 \times 10^{-9}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	$<4.5 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \eta e^+ e^-)/\Gamma_{\text{total}}$	$<1.1 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \eta \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<5.3 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^+ \pi^- e^+ e^-)/\Gamma_{\text{total}}$	$<3.73 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \rho^0 e^+ e^-)/\Gamma_{\text{total}}$	$<1.0 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^+ \pi^- \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<5.5 \times 10^{-7}$ , CL = 90%
$\Gamma(D^0 \rightarrow \rho^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<2.2 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \omega e^+ e^-)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \omega \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<8.3 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow K^- K^+ e^+ e^-)/\Gamma_{\text{total}}$	$<3.15 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \phi e^+ e^-)/\Gamma_{\text{total}}$	$<5.2 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow K^- K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow \phi \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<3.1 \times 10^{-5}$ , CL = 90%
$\Gamma(D^0 \rightarrow K^- \pi^+ e^+ e^-)/\Gamma_{\text{total}}$	$<3.85 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow K^- \pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<3.59 \times 10^{-4}$ , CL = 90%
$\Gamma(D^0 \rightarrow \pi^+ \pi^- \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<8.1 \times 10^{-4}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow K^+ e^+ e^-)/\Gamma_{\text{total}}$	$<3.7 \times 10^{-6}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<2.1 \times 10^{-5}$ , CL = 90%
$\Gamma(D_S^+ \rightarrow K^*(892)^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-3}$ , CL = 90%
$\Gamma(\Lambda_C^+ \rightarrow p e^+ e^-)/\Gamma_{\text{total}}$	$<5.5 \times 10^{-6}$ , CL = 90%
$\Gamma(\Lambda_C^+ \rightarrow p \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<4.4 \times 10^{-5}$ , CL = 90%

## $\Delta B = 1$ WEAK NEUTRAL CURRENT FORBIDDEN

Allowed by higher-order electroweak interactions.

$\Gamma(B^+ \rightarrow \pi^+ \ell^+ \ell^-)/\Gamma_{\text{total}}$	$<4.9 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ e^+ e^-)/\Gamma_{\text{total}}$	$<8.0 \times 10^{-8}$ , CL = 90%
$\Gamma(B^+ \rightarrow \pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(1.79 \pm 0.23) \times 10^{-8}$
$\Gamma(B^+ \rightarrow \pi^+ \nu \bar{\nu})/\Gamma_{\text{total}}$	$<9.8 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ \ell^+ \ell^-)/\Gamma_{\text{total}}$	[Z] $(4.51 \pm 0.23) \times 10^{-7}$ (S = 1.1)
$\Gamma(B^+ \rightarrow K^+ e^+ e^-)/\Gamma_{\text{total}}$	$(5.5 \pm 0.7) \times 10^{-7}$
$\Gamma(B^+ \rightarrow K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(4.43 \pm 0.24) \times 10^{-7}$ (S = 1.2)
$\Gamma(B^+ \rightarrow K^+ \pi^0 \nu)/\Gamma_{\text{total}}$	$<1.6 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow \rho^+ \nu \bar{\nu})/\Gamma_{\text{total}}$	$<2.13 \times 10^{-4}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^*(892)^+ \ell^+ \ell^-)/\Gamma_{\text{total}}$	[Z] $(1.01 \pm 0.11) \times 10^{-6}$ (S = 1.1)
$\Gamma(B^+ \rightarrow K^*(892)^+ e^+ e^-)/\Gamma_{\text{total}}$	$(1.55^{+0.40}_{-0.31}) \times 10^{-6}$
$\Gamma(B^+ \rightarrow K^*(892)^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(9.6 \pm 1.0) \times 10^{-7}$
$\Gamma(B^+ \rightarrow K^*(892)^+ \nu \bar{\nu})/\Gamma_{\text{total}}$	$<4.0 \times 10^{-5}$ , CL = 90%
$\Gamma(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(4.4 \pm 0.4) \times 10^{-7}$
$\Gamma(B^+ \rightarrow \phi K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(7.9^{+2.1}_{-1.7}) \times 10^{-8}$
$\Gamma(B^0 \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-7}$ , CL = 90%
$\Gamma(B^0 \rightarrow e^+ e^-)/\Gamma_{\text{total}}$	$<8.3 \times 10^{-8}$ , CL = 90%
$\Gamma(B^0 \rightarrow e^+ e^- \gamma)/\Gamma_{\text{total}}$	$<1.2 \times 10^{-7}$ , CL = 90%
$\Gamma(B^0 \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(3.9^{+1.6}_{-1.4}) \times 10^{-10}$
$\Gamma(B^0 \rightarrow \mu^+ \mu^- \gamma)/\Gamma_{\text{total}}$	$<1.6 \times 10^{-7}$ , CL = 90%
$\Gamma(B^0 \rightarrow \mu^+ \mu^- \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<5.3 \times 10^{-9}$ , CL = 90%
$\Gamma(B^0 \rightarrow SP, S \rightarrow \mu^+ \mu^-, P \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	[aa] $<5.1 \times 10^{-9}$ , CL = 90%
$\Gamma(B^0 \rightarrow \tau^+ \tau^-)/\Gamma_{\text{total}}$	$<4.1 \times 10^{-3}$ , CL = 90%
$\Gamma(B^0 \rightarrow \pi^0 \ell^+ \ell^-)/\Gamma_{\text{total}}$	$<5.3 \times 10^{-8}$ , CL = 90%
$\Gamma(B^0 \rightarrow \pi^0 e^+ e^-)/\Gamma_{\text{total}}$	$<8.4 \times 10^{-8}$ , CL = 90%
$\Gamma(B^0 \rightarrow \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<6.9 \times 10^{-8}$ , CL = 90%
$\Gamma(B^0 \rightarrow \eta \ell^+ \ell^-)/\Gamma_{\text{total}}$	$<6.4 \times 10^{-8}$ , CL = 90%
$\Gamma(B^0 \rightarrow \eta e^+ e^-)/\Gamma_{\text{total}}$	$<1.08 \times 10^{-7}$ , CL = 90%
$\Gamma(B^0 \rightarrow \eta \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<1.12 \times 10^{-7}$ , CL = 90%
$\Gamma(B^0 \rightarrow \pi^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<6.9 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow K^0 \ell^+ \ell^-)/\Gamma_{\text{total}}$	[Z] $(3.1^{+0.8}_{-0.7}) \times 10^{-7}$
$\Gamma(B^0 \rightarrow K^0 e^+ e^-)/\Gamma_{\text{total}}$	$(1.6^{+1.0}_{-0.8}) \times 10^{-7}$

$\Gamma(B^0 \rightarrow K^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(3.39 \pm 0.34) \times 10^{-7}$
$\Gamma(B^0 \rightarrow K^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<4.9 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow \rho^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<2.08 \times 10^{-4}$ , CL = 90%
$\Gamma(B^0 \rightarrow K^*(892)^0 \ell^+ \ell^-)/\Gamma_{\text{total}}$	[Z] $(9.9^{+1.2}_{-1.1}) \times 10^{-7}$
$\Gamma(B^0 \rightarrow K^*(892)^0 e^+ e^-)/\Gamma_{\text{total}}$	$(1.03^{+0.19}_{-0.17}) \times 10^{-6}$
$\Gamma(B^0 \rightarrow K^*(892)^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(1.02 \pm 0.09) \times 10^{-6}$
$\Gamma(B^0 \rightarrow K^*(892)^0 \nu \bar{\nu})/\Gamma_{\text{total}}$	$<5.5 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow \phi \nu \bar{\nu})/\Gamma_{\text{total}}$	$<1.27 \times 10^{-4}$ , CL = 90%
$\Gamma(B^0 \rightarrow \text{invisible})/\Gamma_{\text{total}}$	$<2.4 \times 10^{-5}$ , CL = 90%
$\Gamma(B^0 \rightarrow \nu \bar{\nu} \gamma)/\Gamma_{\text{total}}$	$<1.7 \times 10^{-5}$ , CL = 90%
$\Gamma(B \rightarrow S e^+ e^-)/\Gamma_{\text{total}}$	$(6.7 \pm 1.7) \times 10^{-6}$ (S = 2.0)
$\Gamma(B \rightarrow S \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(4.3 \pm 1.0) \times 10^{-6}$
$\Gamma(B \rightarrow S \ell^+ \ell^-)/\Gamma_{\text{total}}$	[Z] $(5.8 \pm 1.3) \times 10^{-6}$ (S = 1.8)
$\Gamma(B \rightarrow \pi \ell^+ \ell^-)/\Gamma_{\text{total}}$	$<5.9 \times 10^{-8}$ , CL = 90%
$\Gamma(B \rightarrow \pi e^+ e^-)/\Gamma_{\text{total}}$	$<1.10 \times 10^{-7}$ , CL = 90%
$\Gamma(B \rightarrow \pi \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<5.0 \times 10^{-8}$ , CL = 90%
$\Gamma(B \rightarrow K e^+ e^-)/\Gamma_{\text{total}}$	$(4.4 \pm 0.6) \times 10^{-7}$
$\Gamma(B \rightarrow K^*(892) e^+ e^-)/\Gamma_{\text{total}}$	$(1.19 \pm 0.20) \times 10^{-6}$ (S = 1.2)
$\Gamma(B \rightarrow K \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(4.4 \pm 0.4) \times 10^{-7}$
$\Gamma(B \rightarrow K^*(892) \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(1.06 \pm 0.09) \times 10^{-6}$
$\Gamma(B \rightarrow K \ell^+ \ell^-)/\Gamma_{\text{total}}$	$(4.8 \pm 0.4) \times 10^{-7}$
$\Gamma(B \rightarrow K^*(892) \ell^+ \ell^-)/\Gamma_{\text{total}}$	$(1.05 \pm 0.10) \times 10^{-6}$
$\Gamma(B \rightarrow K \nu \bar{\nu})/\Gamma_{\text{total}}$	$<1.7 \times 10^{-5}$ , CL = 90%
$\Gamma(B \rightarrow K^* \nu \bar{\nu})/\Gamma_{\text{total}}$	$<7.6 \times 10^{-5}$ , CL = 90%
$\Gamma(\bar{B} \rightarrow \bar{S} \nu \nu)/\Gamma_{\text{total}}$	$<6.4 \times 10^{-4}$ , CL = 90%
$\Gamma(\bar{B} \rightarrow e^+ e^- \text{ anything})/\Gamma_{\text{total}}$	—
$\Gamma(\bar{B} \rightarrow \mu^+ \mu^- \text{ anything})/\Gamma_{\text{total}}$	$<3.2 \times 10^{-4}$ , CL = 90%
$\Gamma(\bar{B} \rightarrow \nu \bar{\nu} \text{ anything})/\Gamma_{\text{total}}$	—
$\Gamma(B_S^0 \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$	$<3.1 \times 10^{-6}$ , CL = 90%
$\Gamma(B_S^0 \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(2.9^{+0.7}_{-0.6}) \times 10^{-9}$
$\Gamma(B_S^0 \rightarrow e^+ e^-)/\Gamma_{\text{total}}$	$<2.8 \times 10^{-7}$ , CL = 90%
$\Gamma(B_S^0 \rightarrow \mu^+ \mu^- \mu^+ \mu^-)/\Gamma_{\text{total}}$	$<1.2 \times 10^{-8}$ , CL = 90%
$\Gamma(B_S^0 \rightarrow SP, S \rightarrow \mu^+ \mu^-, P \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$	[aa] $<1.2 \times 10^{-8}$ , CL = 90%
$\Gamma(B_S^0 \rightarrow \phi(1020) \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(8.2 \pm 1.2) \times 10^{-7}$
$\Gamma(B_S^0 \rightarrow \pi^+ \pi^- \mu^+ \mu^-)/\Gamma_{\text{total}}$	$(8.4 \pm 1.7) \times 10^{-8}$
$\Gamma(B_S^0 \rightarrow \phi \nu \bar{\nu})/\Gamma_{\text{total}}$	$<5.4 \times 10^{-3}$ , CL = 90%

## $\Delta T = 1$ WEAK NEUTRAL CURRENT FORBIDDEN

Allowed by higher-order electroweak interactions.

$\Gamma(t \rightarrow Z q(q=u,c))/\Gamma_{\text{total}}$	[bb] $<5 \times 10^{-4}$ , CL = 95%
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### NOTES

In this Summary Table:

When a quantity has “(S = ...)” to its right, the error on the quantity has been enlarged by the “scale factor” S, defined as  $S = \sqrt{\chi^2/(N-1)}$ , where  $N$  is the number of measurements used in calculating the quantity. We do this when  $S > 1$ , which often indicates that the measurements are inconsistent. When  $S > 1.25$ , we also show in the Particle Listings an ideogram of the measurements. For more about S, see the Introduction.

[a] C parity forbids this to occur as a single-photon process.

[b] See the Particle Listings for the (complicated) definition of this quantity.

[c] Time-reversal invariance requires this to be  $0^\circ$  or  $180^\circ$ .

[d] This coefficient is zero if time invariance is not violated.

[e] Allowed by higher-order electroweak interactions.

[f] Violates  $CP$  in leading order. Test of direct  $CP$  violation since the indirect  $CP$ -violating and  $CP$ -conserving contributions are expected to be suppressed.

[g] In the 2010 *Review*, the values for these quantities were given using a measure of the asymmetry that was inconsistent with the usual definition.

[h]  $\text{Re}(\epsilon'/\epsilon) = \epsilon'/\epsilon$  to a very good approximation provided the phases satisfy  $CPT$  invariance.

[i] This mode includes gammas from inner bremsstrahlung but not the direct emission mode  $K_L^0 \rightarrow \pi^+ \pi^- \gamma(\text{DE})$ .

## Tests of Conservation Laws

- [j] Neglecting photon channels. See, e.g., A. Pais and S.B. Treiman, Phys. Rev. **D12**, 2744 (1975).
- [k] Derived from measured values of  $\phi_{+-}$ ,  $\phi_{00}$ ,  $|\eta|$ ,  $|m_{K_L^0} - m_{K_S^0}|$ , and  $\tau_{K_S^0}$ , as described in the introduction to "Tests of Conservation Laws."
- [l] The  $|m_p - m_{\bar{p}}|/m_p$  and  $|q_p + q_{\bar{p}}|/e$  are not independent, and both use the more precise measurement of  $|q_{\bar{p}}/m_{\bar{p}}|/(q_p/m_p)$ .
- [n] The value is for the sum of the charge states or particle/antiparticle states indicated.
- [o] A test of additive vs. multiplicative lepton family number conservation.
- [p] The sign of  $\Delta m_{32}^2$  is not known at this time. The range quoted is for the absolute value.
- [q] Derived from an analysis of neutrino-oscillation experiments.
- [r] This limit is for either  $D^0$  or  $\bar{D}^0$  to  $p e^-$ .
- [s] This limit is for either  $D^0$  or  $\bar{D}^0$  to  $\bar{p} e^+$ .
- [t] The first limit is for  $p \rightarrow$  anything or "disappearance" modes of a bound proton. The second entry, a rough range of limits, assumes the dominant decay modes are among those investigated. For antiprotons the best limit, inferred from the observation of cosmic ray  $\bar{p}$ 's is  $\tau_{\bar{p}} > 10^7$  yr, the cosmic-ray storage time, but this limit depends on a number of assumptions. The best direct observation of stored antiprotons gives  $\tau_{\bar{p}}/B(\bar{p} \rightarrow e^- \gamma) > 7 \times 10^5$  yr.
- [u] There is some controversy about whether nuclear physics and model dependence complicate the analysis for bound neutrons (from which the best limit comes). The first limit here is from reactor experiments with free neutrons.
- [v] This is the best limit for the mode  $e^- \rightarrow \nu \gamma$ . The best limit for "electron disappearance" is  $6.4 \times 10^{24}$  yr.
- [x] See the  $K_S^0$  Particle Listings for the energy limits used in this measurement.
- [y] See the  $K_L^0$  Particle Listings for the energy limits used in this measurement.
- [z] An  $\ell$  indicates an  $e$  or a  $\mu$  mode, not a sum over these modes.
- [aa] Here  $S$  and  $P$  are the hypothetical scalar and pseudoscalar particles with masses of  $2.5 \text{ GeV}/c^2$  and  $214.3 \text{ MeV}/c^2$ , respectively.
- [bb] This limit is for  $\Gamma(t \rightarrow Z q)/\Gamma(t \rightarrow W b)$ .

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## 1. PHYSICAL CONSTANTS

**Table 1.1.** Reviewed 2015 by P.J. Mohr and D.B. Newell (NIST). Mainly from the “CODATA Recommended Values of the Fundamental Physical Constants: 2014” by P.J. Mohr, D.B. Newell, and B.N. Taylor in arXiv:1507.07956 (2015) and RMP (to be submitted). The last group of constants (beginning with the Fermi coupling constant) comes from the Particle Data Group. The figures in parentheses after the values give the 1-standard-deviation uncertainties in the last digits; the corresponding fractional uncertainties in parts per  $10^9$  (ppb) are given in the last column. This set of constants (aside from the last group) is recommended for international use by CODATA (the Committee on Data for Science and Technology). The full 2014 CODATA set of constants may be found at <http://physics.nist.gov/constants>. See also P.J. Mohr and D.B. Newell, “Resource Letter FC-1: The Physics of Fundamental Constants,” Am. J. Phys. **78**, 338 (2010).

Quantity	Symbol, equation	Value	Uncertainty (ppb)
speed of light in vacuum	$c$	299 792 458 m s <sup>-1</sup>	exact*
Planck constant	$h$	6.626 070 040(81)×10 <sup>-34</sup> J s	12
Planck constant, reduced	$\hbar \equiv h/2\pi$	1.054 571 800(13)×10 <sup>-34</sup> J s = 6.582 119 514(40)×10 <sup>-22</sup> MeV s	12 6.1
electron charge magnitude	$e$	1.602 176 6208(98)×10 <sup>-19</sup> C = 4.803 204 673(30)×10 <sup>-10</sup> esu	6.1, 6.1
conversion constant	$\hbar c$	197.326 9788(12) MeV fm	6.1
conversion constant	$(\hbar c)^2$	0.389 379 3656(48) GeV <sup>2</sup> mbarn	12
electron mass	$m_e$	0.510 998 9461(31) MeV/c <sup>2</sup> = 9.109 383 56(11)×10 <sup>-31</sup> kg	6.2, 12
proton mass	$m_p$	938.272 0813(58) MeV/c <sup>2</sup> = 1.672 621 898(21)×10 <sup>-27</sup> kg = 1.007 276 466 879(91) u = 1836.152 673 89(17) $m_e$	6.2, 12 0.090, 0.095
deuteron mass	$m_d$	1875.612 928(12) MeV/c <sup>2</sup>	6.2
unified atomic mass unit (u)	(mass <sup>12</sup> C atom)/12 = (1 g)/(N <sub>A</sub> mol)	931.494 0954(57) MeV/c <sup>2</sup> = 1.660 539 040(20)×10 <sup>-27</sup> kg	6.2, 12
permittivity of free space	$\epsilon_0 = 1/\mu_0 c^2$	8.854 187 817 ... ×10 <sup>-12</sup> F m <sup>-1</sup>	exact
permeability of free space	$\mu_0$	4π × 10 <sup>-7</sup> N A <sup>-2</sup> = 12.566 370 614 ... ×10 <sup>-7</sup> N A <sup>-2</sup>	exact
fine-structure constant	$\alpha = e^2/4\pi\epsilon_0\hbar c$	7.297 352 5664(17)×10 <sup>-3</sup> = 1/137.035 999 139(31) <sup>†</sup>	0.23, 0.23
classical electron radius	$r_e = e^2/4\pi\epsilon_0 m_e c^2$	2.817 940 3227(19)×10 <sup>-15</sup> m	0.68
(e <sup>-</sup> Compton wavelength)/2π	$\lambda_e = \hbar/m_e c = r_e \alpha^{-1}$	3.861 592 6764(18)×10 <sup>-13</sup> m	0.45
Bohr radius ( $m_{\text{nucleus}} = \infty$ )	$a_\infty = 4\pi\epsilon_0\hbar^2/m_e e^2 = r_e \alpha^{-2}$	0.529 177 210 67(12)×10 <sup>-10</sup> m	0.23
wavelength of 1 eV/c particle	$\hbar c/(1 \text{ eV})$	1.239 841 9739(76)×10 <sup>-6</sup> m	6.1
Rydberg energy	$\hbar c R_\infty = m_e e^4/2(4\pi\epsilon_0)^2 \hbar^2 = m_e c^2 \alpha^2/2$	13.605 693 009(84) eV	6.1
Thomson cross section	$\sigma_T = 8\pi r_e^2/3$	0.665 245 871 58(91) barn	1.4
Bohr magneton	$\mu_B = e\hbar/2m_e$	5.788 381 8012(26)×10 <sup>-11</sup> MeV T <sup>-1</sup>	0.45
nuclear magneton	$\mu_N = e\hbar/2m_p$	3.152 451 2550(15)×10 <sup>-14</sup> MeV T <sup>-1</sup>	0.46
electron cyclotron freq./field	$\omega_{\text{cycl}}^e/B = e/m_e$	1.758 820 024(11)×10 <sup>11</sup> rad s <sup>-1</sup> T <sup>-1</sup>	6.2
proton cyclotron freq./field	$\omega_{\text{cycl}}^p/B = e/m_p$	9.578 833 226(59)×10 <sup>7</sup> rad s <sup>-1</sup> T <sup>-1</sup>	6.2
gravitational constant <sup>‡</sup>	$G_N$	6.674 08(31)×10 <sup>-11</sup> m <sup>3</sup> kg <sup>-1</sup> s <sup>-2</sup> = 6.708 61(31)×10 <sup>-39</sup> $\hbar c$ (GeV/c <sup>2</sup> ) <sup>-2</sup>	4.7 × 10 <sup>4</sup> 4.7 × 10 <sup>4</sup>
standard gravitational accel.	$g_N$	9.806 65 m s <sup>-2</sup>	exact
Avogadro constant	$N_A$	6.022 140 857(74)×10 <sup>23</sup> mol <sup>-1</sup>	12
Boltzmann constant	$k$	1.380 648 52(79)×10 <sup>-23</sup> J K <sup>-1</sup> = 8.617 3303(50)×10 <sup>-5</sup> eV K <sup>-1</sup>	570 570
molar volume, ideal gas at STP	$N_A k(273.15 \text{ K})/(101\,325 \text{ Pa})$	22.413 962(13)×10 <sup>-3</sup> m <sup>3</sup> mol <sup>-1</sup>	570
Wien displacement law constant	$b = \lambda_{\text{max}} T$	2.897 7729(17)×10 <sup>-3</sup> m K	570
Stefan-Boltzmann constant	$\sigma = \pi^2 k^4/60\hbar^3 c^2$	5.670 367(13)×10 <sup>-8</sup> W m <sup>-2</sup> K <sup>-4</sup>	2300
Fermi coupling constant**	$G_F/(\hbar c)^3$	1.166 378 7(6)×10 <sup>-5</sup> GeV <sup>-2</sup>	500
weak-mixing angle	$\sin^2 \hat{\theta}(M_Z) \text{ (}\overline{\text{MS}}\text{)}$	0.231 29(5) <sup>††</sup>	2.2 × 10 <sup>5</sup>
$W^\pm$ boson mass	$m_W$	80.385(15) GeV/c <sup>2</sup>	1.9 × 10 <sup>5</sup>
$Z^0$ boson mass	$m_Z$	91.1876(21) GeV/c <sup>2</sup>	2.3 × 10 <sup>4</sup>
strong coupling constant	$\alpha_s(m_Z)$	0.1182(12)	1.0 × 10 <sup>7</sup>
$\pi = 3.141\,592\,653\,589\,793\,238$ $e = 2.718\,281\,828\,459\,045\,235$ $\gamma = 0.577\,215\,664\,901\,532\,861$			
1 in ≡ 0.0254 m	1 G ≡ 10 <sup>-4</sup> T	1 eV = 1.602 176 6208(98) × 10 <sup>-19</sup> J	$kT$ at 300 K = [38.681 740(22)] <sup>-1</sup> eV
1 Å ≡ 0.1 nm	1 dyne ≡ 10 <sup>-5</sup> N	1 eV/c <sup>2</sup> = 1.782 661 907(11) × 10 <sup>-36</sup> kg	0 °C ≡ 273.15 K
1 barn ≡ 10 <sup>-28</sup> m <sup>2</sup>	1 erg ≡ 10 <sup>-7</sup> J	2.997 924 58 × 10 <sup>9</sup> esu = 1 C	1 atmosphere ≡ 760 Torr ≡ 101 325 Pa

\* The meter is the length of the path traveled by light in vacuum during a time interval of 1/299 792 458 of a second.

† At  $Q^2 = 0$ . At  $Q^2 \approx m_W^2$  the value is  $\sim 1/128$ .

‡ Absolute lab measurements of  $G_N$  have been made only on scales of about 1 cm to 1 m.

\*\* See the discussion in Sec. 10, “Electroweak model and constraints on new physics.”

†† The corresponding  $\sin^2 \theta$  for the effective angle is 0.23155(5).

## 2. ASTROPHYSICAL CONSTANTS AND PARAMETERS

**Table 2.1.** Revised March 2016 by D.E. Groom (LBNL). The figures in parentheses after some values give the  $1\text{-}\sigma$  uncertainties in the last digit(s). Physical constants are from Ref. 1. While every effort has been made to obtain the most accurate current values of the listed quantities, the table does not represent a critical review or adjustment of the constants, and is not intended as a primary reference.

The values and uncertainties for the cosmological parameters depend on the exact data sets, priors, and basis parameters used in the fit. Many of the derived parameters reported in this table have non-Gaussian likelihoods. Parameters may be highly correlated, so care must be taken in propagating errors. Unless otherwise specified, cosmological parameters are derived from 6-parameter fits to a flat  $\Lambda$ CDM cosmology *Planck* 2015 temperature (TT) + low  $\ell$  polarization data (lowP) + lensing [2]. For more information see Ref. 3 and the original papers.

Quantity	Symbol, equation	Value	Reference, footnote
speed of light	$c$	$299\,792\,458\text{ m s}^{-1}$	exact[4]
Newtonian constant of gravitation	$G_N$	$6.674\,08(31) \times 10^{-11}\text{ m}^3\text{ kg}^{-1}\text{ s}^{-2}$	[1]
Planck mass	$\sqrt{\hbar c/G_N}$	$1.220\,910(29) \times 10^{19}\text{ GeV}/c^2 = 2.176\,47(5) \times 10^{-8}\text{ kg}$	[1]
Planck length	$\sqrt{\hbar G_N/c^3}$	$1.616\,229(38) \times 10^{-35}\text{ m}$	[1]
standard acceleration of gravity	$g_N$	$9.806\,65\text{ m s}^{-2}$	exact[1]
jansky (flux density)	Jy	$10^{-26}\text{ W m}^{-2}\text{ Hz}^{-1}$	definition
tropical year (equinox to equinox) (2011)	yr	$31\,556\,925.2\text{ s} \approx \pi \times 10^7\text{ s}$	[5]
sidereal year (fixed star to fixed star) (2011)		$31\,558\,149.8\text{ s} \approx \pi \times 10^7\text{ s}$	[5]
mean sidereal day (2011) (time between vernal equinox transits)		$23^{\text{h}}\,56^{\text{m}}\,04^{\text{s}}.090\,53$	[5]
astronomical unit	au	$149\,597\,870\,700\text{ m}$	exact[6]
parsec (1 au/1 arc sec)	pc	$3.085\,677\,581\,49 \times 10^{16}\text{ m} = 3.262 \dots\text{ ly}$	exact[7]
light year (deprecated unit)	ly	$0.306\,6 \dots\text{ pc} = 0.946\,053 \dots \times 10^{16}\text{ m}$	
Schwarzschild radius of the Sun	$2G_N M_\odot/c^2$	$2.953\,250\,24\text{ km}$	[8]
Solar mass	$M_\odot$	$1.988\,48(9) \times 10^{30}\text{ kg}$	[9]
nominal Solar equatorial radius	$R_\odot$	$6.957 \times 10^8\text{ m}$	exact[10]
nominal Solar constant	$S_\odot$	$1361\text{ W m}^{-2}$	exact[10,11]
nominal Solar photosphere temperature	$T_\odot$	$5772\text{ K}$	exact[10]
nominal Solar luminosity	$L_\odot$	$3.828 \times 10^{26}\text{ W}$	exact[10,12]
Schwarzschild radius of the Earth	$2G_N M_\oplus/c^2$	$8.870\,056\,580(18)\text{ mm}$	[13]
Earth mass	$M_\oplus$	$5.972\,4(3) \times 10^{24}\text{ kg}$	[14]
nominal Earth equatorial radius	$R_\oplus$	$6.3781 \times 10^6\text{ m}$	exact[10]
luminosity conversion	$L$	$3.0128 \times 10^{28} \times 10^{-0.4 M_{\text{bol}}}\text{ W}$	[15]
flux conversion	$\mathcal{F}$	$(M_{\text{bol}} = \text{absolute bolometric magnitude} = \text{bolometric magnitude at } 10\text{ pc})$ $2.5180 \times 10^{-8} \times 10^{-0.4 m_{\text{bol}}}\text{ W m}^{-2}$	[15]
ABsolute monochromatic magnitude	AB	$(m_{\text{bol}} = \text{apparent bolometric magnitude})$ $-2.5 \log_{10} f_\nu - 56.10\text{ (for } f_\nu \text{ in W m}^{-2}\text{ Hz}^{-1}\text{)}$ $= -2.5 \log_{10} f_\nu + 8.90\text{ (for } f_\nu \text{ in Jy)}$	[16]
Solar angular velocity around the Galactic center	$\Theta_0/R_0$	$30.3 \pm 0.9\text{ km s}^{-1}\text{ kpc}^{-1}$	[17]
Solar distance from Galactic center	$R_0$	$8.00 \pm 0.25\text{ kpc}$	[17,18]
circular velocity at $R_0$	$v_0$ or $\Theta_0$	$254(16)\text{ km s}^{-1}$	[17]
escape velocity from Galaxy	$v_{\text{esc}}$	$498\text{ km/s} < v_{\text{esc}} < 608\text{ km/s}$	[19]
local disk density	$\rho_{\text{disk}}$	$3\text{--}12 \times 10^{-24}\text{ g cm}^{-3} \approx 2\text{--}7\text{ GeV}/c^2\text{ cm}^{-3}$	[20]
local dark matter density	$\rho_\chi$	canonical value $0.3\text{ GeV}/c^2\text{ cm}^{-3}$ within factor 2-3	[21]
present day CMB temperature	$T_0$	$2.7255(6)\text{ K}$	[22,24]
present day CMB dipole amplitude		$3.3645(20)\text{ mK}$	[22,23]
Solar velocity with respect to CMB		$369(1)\text{ km s}^{-1}$ towards $(\ell, b) = (263.99(14)^\circ, 48.26(3)^\circ)$	[22,25]
Local Group velocity with respect to CMB	$v_{\text{LG}}$	$627(22)\text{ km s}^{-1}$ towards $(\ell, b) = (276(3)^\circ, 30(3)^\circ)$	[22,25]
number density of CMB photons	$n_\gamma$	$410.7(T/2.7255)^3\text{ cm}^{-3}$	[26]
density of CMB photons	$\rho_\gamma$	$4.645(4)(T/2.7255)^4 \times 10^{-34}\text{ g cm}^{-3} \approx 0.260\text{ eV cm}^{-3}$	[26]
entropy density/Boltzmann constant	$s/k$	$2.891.2(T/2.7255)^3\text{ cm}^{-3}$	[26]
present day Hubble expansion rate	$H_0$	$100\text{ h km s}^{-1}\text{ Mpc}^{-1} = h \times (9.777\,752\text{ Gyr})^{-1}$	[27]
scale factor for Hubble expansion rate	$h$	$0.678(9)$	[2,3]
Hubble length	$c/H_0$	$0.925\,0629 \times 10^{26}\text{ h}^{-1}\text{ m} = 1.374(18) \times 10^{26}\text{ m}$	
scale factor for cosmological constant	$c^2/3H_0^2$	$2.85247 \times 10^{51}\text{ h}^{-2}\text{ m}^2 = 6.20(17) \times 10^{51}\text{ m}^2$	
critical density of the Universe	$\rho_{\text{crit}} = 3H_0^2/8\pi G_N$	$1.878\,40(9) \times 10^{-29}\text{ h}^2\text{ g cm}^{-3}$ $= 1.053\,71(5) \times 10^{-5}\text{ h}^2\text{ (GeV}/c^2\text{) cm}^{-3}$ $= 2.775\,37(13) \times 10^{11}\text{ h}^2\text{ M}_\odot\text{Mpc}^{-3}$	
baryon-to-photon ratio (from BBN)	$\eta = n_b/n_\gamma$	$5.8 \times 10^{-10} \leq \eta \leq 6.6 \times 10^{-10}\text{ (95\% CL)}$	[28]
number density of baryons	$n_b$	$2.503(26) \times 10^{-7}\text{ cm}^{-3}$ $(2.4 \times 10^{-7} < n_b < 2.7 \times 10^{-7})\text{ cm}^{-3}\text{ (95\% CL)}$	[2,3,29,30] $\eta \times n_\gamma$
CMB radiation density of the Universe	$\Omega_\gamma = \rho_\gamma/\rho_{\text{crit}}$	$2.473 \times 10^{-5}(T/2.7255)^4\text{ h}^{-2} = 5.38(15) \times 10^{-5}$	[26]
--- <i>Planck</i> 2015 6-parameter fit to flat $\Lambda$ CDM cosmology ---			
baryon density of the Universe	$\Omega_b = \rho_b/\rho_{\text{crit}}$	$^{\dagger}0.02226(23)\text{ h}^{-2} = ^{\dagger}0.0484(10)$	[2,3,23]
cold dark matter density of the universe	$\Omega_{\text{CDM}} = \rho_{\text{CDM}}/\rho_{\text{crit}}$	$^{\dagger}0.1186(20)\text{ h}^{-2} = ^{\dagger}0.258(11)$	[2,3,23]
$100 \times \text{approx to } r_*/D_A$	$100 \times \theta_{\text{MC}}$	$^{\dagger}1.0410(5)$	[2,3]
reionization optical depth	$\tau$	$^{\dagger}0.066(16)$	[2,3]
scalar spectral index	$n_s$	$^{\dagger}0.968(6)$	[2,3]
ln pwr primordial curvature pert. ( $k_0=0.05\text{ Mpc}^{-1}$ )	$\ln(10^{10}\Delta_{\mathcal{R}}^2)$	$^{\dagger}3.062(29)$	[2,3]



Quantity	Symbol, equation	Value	Reference, footnote
dark energy density of the $\Lambda$ CDM Universe	$\Omega_\Lambda$	$^\dagger 0.692 \pm 0.012$	[2,3]
pressureless matter density of the Universe	$\Omega_m = \Omega_{\text{CDM}} + \Omega_b$	$^\dagger 0.308 \pm 0.012$	[2,3]
fluctuation amplitude at $8 h^{-1}$ Mpc scale	$\sigma_8$	$^\dagger 0.815 \pm 0.009$	[2,3]
redshift of matter-radiation equality	$z_{\text{eq}}$	$^\dagger 3365 \pm 44$	[2]
redshift at which optical depth equals unity	$z_*$	$^\dagger 1089.9 \pm 0.4$	[2]
comoving size of sound horizon at $z_*$	$r_*$	$^\dagger 144.9 \pm 0.4$ Mpc	(Planck CMB) [31]
age when optical depth equals unity	$t_*$	373 kyr	[32]
redshift at half reionization	$z_{\text{reion}}$	$^\dagger 8.8^{+1.7}_{-1.4}$	[2]
redshift when acceleration was zero	$z_q$	$\sim 0.65$	[32]
age of the Universe	$t_0$	$^\dagger 13.80 \pm 0.04$ Gyr	[2]
effective number of neutrinos	$N_{\text{eff}}$	$^\# 3.1 \pm 0.6$	[2,33]
sum of neutrino masses	$\sum m_\nu$	$^\# < 0.68$ eV (Planck CMB); $\geq 0.05$ eV (mixing)	[2,34,35]
neutrino density of the Universe	$\Omega_\nu = h^{-2} \sum m_{\nu_j} / 93.04 \text{ eV}$	$^\# < 0.016$ (Planck CMB); $\geq 0.0012$ (mixing)	[2,34,35]
curvature	$\Omega_K$	$^\# -0.005^{+0.016}_{-0.017}$ (95%CL)	[2]
running spectral index slope, $k_0 = 0.002 \text{ Mpc}^{-1}$	$dn_s/d \ln k$	$^\# -0.003(15)$	[2]
tensor-to-scalar field perturbations ratio, $k_0 = 0.002 \text{ Mpc}^{-1}$	$r_{0.002} = T/S$	$^\# < 0.114$ at 95% CL; no running	[2,3]
dark energy equation of state parameter	$w$	$-0.97 \pm 0.05$	[31,36]
primordial helium fraction	$Y_p$	$0.245 \pm 0.004$	[22,37]

$^\ddagger$  Parameter in 6-parameter  $\Lambda$ CDM fit [2].

$^\dagger$  Derived parameter in 6-parameter  $\Lambda$ CDM fit [2].

$^\#$  Extended model parameter (TT + lensing) [2].

## References:

- CODATA recommended 2014 values of the fundamental physical constants: [physics.nist.gov/constants](http://physics.nist.gov/constants).
- Planck Collab. 2015 Results XIII, *Astron. & Astrophys.* submitted, [arXiv:1502.01589v2](https://arxiv.org/abs/1502.01589v2).
- O. Lahav & A.R. Liddle, “The Cosmological Parameters,” Sec. 25 in this *Review*.
- B.W. Petley, *Nature* **303**, 373 (1983).
- The Astronomical Almanac Online for the year 2016*, [asa.usno.navy.mil/SecK/Constants.html](http://asa.usno.navy.mil/SecK/Constants.html).
- The astronomical unit of length (the au) in meters is re-defined (resolution B2, IAU XXVIII GA 2012) to be a conventional unit of length in agreement with the value adopted in the IAU 2009 Resolution B2; it is to be used with all time scales.
- The distance at which 1 au subtends 1 arc sec: 1 au divided by  $\pi/648000$ .
- Product of  $2/c^2$  and the observationally determined Solar mass parameter  $G_N M_\odot$  [5]. Truncated to 8 places so that TCB and TDB time scale values agree.
- $G_N M_\odot$  [5]  $\div G_N$  [1].
- XXIXth IAU General Assembly, Resolution B3, “on recommended nominal conversion constants ...” Calligraphic symbol indicates recommended nominal value.
- See also G. Kopp & J.L. Lean, *Geophys. Res. Lett.* **38**, L01706 (2011), who give  $1360.8 \pm 0.6 \text{ W m}^{-2}$ . See paper for caveats and other measurements.
- $4\pi (1 \text{ au})^2 \times \mathcal{S}_\odot$ , assuming isotropic irradiance.
- Product of  $2/c^2$  and the geocentric gravitational constant  $G_N M_\oplus$  [5]. Truncated to 8 places so that TCB, TT, and TDB time scale values agree.
- $G_N M_\oplus$  [5]  $\div G_N$  [1].
- XXIXth IAU General Assembly, Resolution B2, “on recommended zero points for the absolute and apparent bolometric magnitude scales”.
- J. B. Oke and J. E. Gunn, *Astrophys. J.* **266**, 713 (1983). Note that in the definition of AB the sign of the constant is wrong.
- M.J. Reid, *et al.*, *Astrophys. J.* **700**, 137 (2009). Note that  $\Theta_0/R_0$  is better determined than either  $\Theta_0$  or  $R_0$ .
- Z.M. Malkin, [arXiv:1202.6128](https://arxiv.org/abs/1202.6128) and *Astron. Rep.* **57**, 128 (2013). 52 determinations of  $R_0$  over 20 years are given. The weighted mean of these *unevaluated* results is  $7.94 \pm 0.05$  kpc, with  $\chi^2/N_{\text{dof}} = 1.26$ . If the 8 values more than  $3\sigma$  from the mean are eliminated,  $\langle R_0 \rangle = 8.02 \pm 0.06$  kpc and  $\chi^2/N_{\text{dof}} = 0.67$ . The author suggests using  $R_0 = 8.00 \pm 0.25$  kpc.
- M. C. Smith *et al.*, *Mon. Not. R. Astr. Soc.* **379**, 755 (2007).
- G. Gilmore, R.F.G. Wyse, & K. Kuijken, *Ann. Rev. Astron. Astrophys.* **27**, 555 (1989).
- Sampling of many references:  
M. Mori *et al.*, *Phys. Lett.* **B289**, 463 (1992); E.I. Gates *et al.*, *Astrophys. J.* **449**, L133 (1995); M. Kamionkowski & A. Kinkhabwala, *Phys. Rev.* **D57**, 325 (1998); M. Weber & W. de Boer, *Astron. & Astrophys.* **509**, A25 (2010); P. Salucci *et al.*, *Astron. & Astrophys.* **523**, A83 (2010); R. Catena & P. Ullio, *JCAP* **1008**, 004 (2010) conclude  $\rho_{\text{DM}}^{\text{local}} = 0.39 \pm 0.03 \text{ GeV cm}^{-3}$ .
- D. Scott & G.F. Smoot, “Cosmic Microwave Background,” Sec. 28 in this *Review*.
- Planck Collab. 2015 Results I, *Astron. & Astrophys.* submitted, [arXiv:1502.01581v3](https://arxiv.org/abs/1502.01581v3).
- D. Fixsen, *Astrophys. J.* **707**, 916 (2009).
- G. Hinshaw *et al.*, *Astrophys. J. Suppl.* **208**, 19 (2013), [arXiv:1212.5226](https://arxiv.org/abs/1212.5226); D.J. Fixsen *et al.*, *Astrophys. J.* **473**, 576 (1996); A. Kogut *et al.*, *Astrophys. J.* **419**, 1 (1993).
- $n_\gamma = \frac{2\zeta(3)}{\pi^2} \left(\frac{kT}{hc}\right)^3$ ;  $\rho_\gamma = \frac{\pi^2 kT}{15 c^2} \left(\frac{kT}{hc}\right)^3$ ;  $s/k = \frac{2 \cdot 43 \cdot \pi^2}{11 \cdot 45} \left(\frac{kT}{hc}\right)^3$ ;  $kT/hc = 11.902(3)(T/2.7255)/\text{cm}$ .
- Conversion using length of sidereal year.
- B.D. Fields, P. Molarto, & S. Sarkar, “Big-Bang Nucleosynthesis,” in this *Review*.
- $n_b$  depends only upon the measured  $\Omega_b h^2$ , the average baryon mass at the present epoch [30], and  $G_N$ :  $n_b = (\Omega_b h^2)(h^{-2} \rho_{\text{crit}})/(0.93711 \text{ GeV}/c^2 \text{ per baryon})$ .
- G. Steigman, *JCAP* **10**, 016, (2006).
- D.H. Weinberg & M. White, “Dark Energy,” Sec. 27 in this *Review*.
- D. Scott, A. Narimani, & D.N. Page, [arXiv:1309.2381v2](https://arxiv.org/abs/1309.2381v2).
- Summary Tables in this *Review* list  $N_\nu = 2.984(8)$  (Standard Model fits to LEP-SLC data). Because neutrinos are not completely decoupled at  $e^\pm$  annihilation, the effective number of massless neutrino species is 3.046, rather than 3.
- The sum is over all neutrino mass eigenstates. The lower limit follows from neutrino mixing results reported in this *Review* combined with the assumptions that there are three light neutrinos ( $m_\nu < 45 \text{ GeV}/c^2$ ) and that the lightest neutrino is substantially less massive than the others:  $\Delta m_{32}^2 = (2.44 \pm 0.06) \times 10^{-3} \text{ eV}^2$ , so  $\sum m_{\nu_j} \geq m_{\nu_3} \approx \sqrt{\Delta m_{32}^2} = 0.05 \text{ eV}$ . About the same limit obtains if the mass hierarchy is inverted, with  $m_{\nu_1} \approx m_{\nu_2} \gg m_{\nu_3}$ . Alternatively, if the limit obtained from tritium decay experiments ( $m_\nu < 2 \text{ eV}$ ) is used for the upper limit, then  $\Omega_\nu < 0.05$ .
- Astrophysical determinations of  $\sum m_{\nu_j}$ , reported in the Full Listings of this *Review* under “Sum of the neutrino masses,” range from  $< 0.17 \text{ eV}$  to  $< 2.3 \text{ eV}$  in papers published since 2003.
- É. Auborg *et al.*, *Phys. Rev.* **D92**, 123516 (2015).
- E. Aver *et al.*, *JCAP* **07**, 011 (2015).

3. INTERNATIONAL SYSTEM OF UNITS (SI)

See “The International System of Units (SI),” NIST Special Publication **330**, B.N. Taylor, ed. (USGPO, Washington, DC, 1991); and “Guide for the Use of the International System of Units (SI),” NIST Special Publication **811**, 1995 edition, B.N. Taylor (USGPO, Washington, DC, 1995).

Physical quantity	Name of unit	Symbol
<i>Base units</i>		
length	meter	m
mass	kilogram	kg
time	second	s
electric current	ampere	A
thermodynamic temperature	kelvin	K
amount of substance	mole	mol
luminous intensity	candela	cd
<i>Derived units with special names</i>		
plane angle	radian	rad
solid angle	steradian	sr
frequency	hertz	Hz
energy	joule	J
force	newton	N
pressure	pascal	Pa
power	watt	W
electric charge	coulomb	C
electric potential	volt	V
electric resistance	ohm	$\Omega$
electric conductance	siemens	S
electric capacitance	farad	F
magnetic flux	weber	Wb
inductance	henry	H
magnetic flux density	tesla	T
luminous flux	lumen	lm
illuminance	lux	lx
celsius temperature	degree celsius	$^{\circ}\text{C}$
activity (of a radioactive source)*	becquerel	Bq
absorbed dose (of ionizing radiation)*	gray	Gy
dose equivalent*	sievert	Sv

SI prefixes		
$10^{24}$	yotta	(Y)
$10^{21}$	zetta	(Z)
$10^{18}$	exa	(E)
$10^{15}$	peta	(P)
$10^{12}$	tera	(T)
$10^9$	giga	(G)
$10^6$	mega	(M)
$10^3$	kilo	(k)
$10^2$	hecto	(h)
10	deca	(da)
$10^{-1}$	deci	(d)
$10^{-2}$	centi	(c)
$10^{-3}$	milli	(m)
$10^{-6}$	micro	( $\mu$ )
$10^{-9}$	nano	(n)
$10^{-12}$	pico	(p)
$10^{-15}$	femto	(f)
$10^{-18}$	atto	(a)
$10^{-21}$	zepto	(z)
$10^{-24}$	yocto	(y)

\*See our section 36, on “Radioactivity and radiation protection,” p. 494.

**Table 4.1.** Revised June 2016 by D.E. Groom (LBNL). The atomic number (top left) is the number of protons in the nucleus. The atomic masses (bottom) of stable elements are weighted by isotopic abundances in the Earth's surface. Atomic masses are relative to the mass of  $^{12}\text{C}$ , defined to be exactly 12 unified atomic mass units (u) (approx. g/mole). The exceptions are Th, Pa, and U, which have no stable isotopes but do have characteristic terrestrial compositions. Relative isotopic abundances often vary considerably, both in natural and commercial samples; this is reflected in the number of significant figures given for the mass. Masses may be found at [http://physics.nist.gov/cgi-bin/Compositions/stand\\_alone.pl](http://physics.nist.gov/cgi-bin/Compositions/stand_alone.pl). If there is no stable isotope the atomic mass of the most stable isotope is given in parentheses.

IUPAC announced verification of the discoveries of elements 113, 115, 117, and 118 in December 2015. Provisional names were assigned in June 2016. The 7th period of the periodic table is now complete.

1 IA																	18 VIIIA										
1 H hydrogen 1.008	2 IIA		PERIODIC TABLE OF THE ELEMENTS														13 IIIA		14 IVA		15 VA		16 VIA		17 VIIA		2 He helium 4.002602
3 Li lithium 6.94	4 Be beryllium 9.012182	5 B boron 10.81															6 C carbon 12.0107	7 N nitrogen 14.007	8 O oxygen 15.999	9 F fluorine 18.998403163	10 Ne neon 20.1797						
11 Na sodium 22.98976928	12 Mg magnesium 24.305	13 Al aluminum 26.9815385															14 Si silicon 28.085	15 P phosphorus 30.973761998	16 S sulfur 32.06	17 Cl chlorine 35.45	18 Ar argon 39.948						
3 IIIB		4 IVB															5 VB		6 VIB		7 VIIB		8 VIII		9 IX		10 X
19 K potassium 39.0983	20 Ca calcium 40.078	21 Sc scandium 44.955908	22 Ti titanium 47.867	23 V vanadium 50.9415	24 Cr chromium 51.9961	25 Mn manganese 54.938044	26 Fe iron 55.845	27 Co cobalt 58.933195	28 Ni nickel 58.6934	29 Cu copper 63.546	30 Zn zinc 65.38	31 Ga gallium 69.723	32 Ge germanium 72.630	33 As arsenic 74.921595	34 Se selenium 78.971	35 Br bromine 79.904	36 Kr krypton 83.798										
37 Rb rubidium 85.4678	38 Sr strontium 87.62	39 Y yttrium 88.90584	40 Zr zirconium 91.224	41 Nb niobium 92.90637	42 Mo molybdenum 95.95	43 Tc technetium (97.907212)	44 Ru ruthenium 101.07	45 Rh rhodium 102.90550	46 Pd palladium 106.42	47 Ag silver 107.8682	48 Cd cadmium 112.414	49 In indium 114.818	50 Sn tin 118.710	51 Sb antimony 121.760	52 Te tellurium 127.60	53 I iodine 126.90447	54 Xe xenon 131.293										
55 Cs caesium 132.90545196	56 Ba barium 137.327	57-71 LANTH- ANIDES	72 Hf hafnium 178.49	73 Ta tantalum 180.94788	74 W tungsten 183.84	75 Re rhenium 186.207	76 Os osmium 190.23	77 Ir iridium 192.217	78 Pt platinum 195.084	79 Au gold 196.966569	80 Hg mercury 200.592	81 Tl thallium 204.38	82 Pb lead 207.2	83 Bi bismuth 208.98040	84 Po polonium (208.98243)	85 At astatine (209.98715)	86 Rn radon (222.01758)										
87 Fr francium (223.01974)	88 Ra radium (226.02541)	89-103 ACTINIDES	104 Rf rutherford. (267.12169)	105 Db dubnium (268.12567)	106 Sg seaborgium (271.13393)	107 Bh bohrium (272.13826)	108 Hs hassium (270.13429)	109 Mt meitnerium (276.15159)	110 Ds darmstadt. (281.16451)	111 Rg roentgen. (280.16514)	112 Cn copernicium (285.17712)	113 Nh (nihonium) (284.17873)	114 Fl flerovium (289.19042)	115 Mc (moscovium) (288.19274)	116 Lv livermorium (293.20449)	117 Ts (tennessine) (292.20746)	118 Og (oganeson) (294.21392)										

Lanthanide series	57 La lanthanum 138.90547	58 Ce cerium 140.116	59 Pr praseodym. 140.90766	60 Nd neodymium 144.242	61 Pm promethium (144.91276)	62 Sm samarium 150.36	63 Eu europium 151.964	64 Gd gadolinium 157.25	65 Tb terbium 158.92535	66 Dy dysprosium 162.500	67 Ho holmium 164.93033	68 Er erbium 167.259	69 Tm thulium 168.93422	70 Yb ytterbium 173.054	71 Lu lutetium 174.9668
Actinide series	89 Ac actinium (227.02775)	90 Th thorium 232.0377	91 Pa protactinium 231.03588	92 U uranium 238.02891	93 Np neptunium (237.04817)	94 Pu plutonium (244.06420)	95 Am americium (243.06138)	96 Cm curium (247.07035)	97 Bk berkelium (247.07031)	98 Cf californium (251.07959)	99 Es einsteinium (252.08298)	100 Fm fermium (257.09511)	101 Md mendelevium (258.09844)	102 No nobelium (259.10103)	103 Lr lawrencium (262.10961)

## 5. ELECTRONIC STRUCTURE OF THE ELEMENTS

**Table 5.1.** Reviewed 2011 by J.E. Sansonetti (NIST). The electronic configurations and the ionization energies are from the NIST database, “Ground Levels and Ionization Energies for the Neutral Atoms,” W.C. Martin, A. Musgrove, S. Kotochigova, and J.E. Sansonetti, [http://www.nist.gov/pml/data/ion\\_energy.cfm](http://www.nist.gov/pml/data/ion_energy.cfm). The electron configuration for, say, iron indicates an argon electronic core (see argon) plus six 3d electrons and two 4s electrons.

Element			Electron configuration ( $3d^5$ = five 3d electrons, <i>etc.</i> )	Ground state $2S+1L_J$	Ionization energy (eV)
1	H	Hydrogen	1s	$^2S_{1/2}$	13.5984
2	He	Helium	1s <sup>2</sup>	$^1S_0$	24.5874
3	Li	Lithium	(He) 2s	$^2S_{1/2}$	5.3917
4	Be	Beryllium	(He) 2s <sup>2</sup>	$^1S_0$	9.3227
5	B	Boron	(He) 2s <sup>2</sup> 2p	$^2P_{1/2}$	8.2980
6	C	Carbon	(He) 2s <sup>2</sup> 2p <sup>2</sup>	$^3P_0$	11.2603
7	N	Nitrogen	(He) 2s <sup>2</sup> 2p <sup>3</sup>	$^4S_{3/2}$	14.5341
8	O	Oxygen	(He) 2s <sup>2</sup> 2p <sup>4</sup>	$^3P_2$	13.6181
9	F	Fluorine	(He) 2s <sup>2</sup> 2p <sup>5</sup>	$^2P_{3/2}$	17.4228
10	Ne	Neon	(He) 2s <sup>2</sup> 2p <sup>6</sup>	$^1S_0$	21.5645
11	Na	Sodium	(Ne) 3s	$^2S_{1/2}$	5.1391
12	Mg	Magnesium	(Ne) 3s <sup>2</sup>	$^1S_0$	7.6462
13	Al	Aluminum	(Ne) 3s <sup>2</sup> 3p	$^2P_{1/2}$	5.9858
14	Si	Silicon	(Ne) 3s <sup>2</sup> 3p <sup>2</sup>	$^3P_0$	8.1517
15	P	Phosphorus	(Ne) 3s <sup>2</sup> 3p <sup>3</sup>	$^4S_{3/2}$	10.4867
16	S	Sulfur	(Ne) 3s <sup>2</sup> 3p <sup>4</sup>	$^3P_2$	10.3600
17	Cl	Chlorine	(Ne) 3s <sup>2</sup> 3p <sup>5</sup>	$^2P_{3/2}$	12.9676
18	Ar	Argon	(Ne) 3s <sup>2</sup> 3p <sup>6</sup>	$^1S_0$	15.7596
19	K	Potassium	(Ar) 4s	$^2S_{1/2}$	4.3407
20	Ca	Calcium	(Ar) 4s <sup>2</sup>	$^1S_0$	6.1132
21	Sc	Scandium	(Ar) 3d 4s <sup>2</sup>	$^2D_{3/2}$	6.5615
22	Ti	Titanium	(Ar) 3d <sup>2</sup> 4s <sup>2</sup>	$^3F_2$	6.8281
23	V	Vanadium	(Ar) 3d <sup>3</sup> 4s <sup>2</sup>	$^4F_{3/2}$	6.7462
24	Cr	Chromium	(Ar) 3d <sup>5</sup> 4s	$^7S_3$	6.7665
25	Mn	Manganese	(Ar) 3d <sup>5</sup> 4s <sup>2</sup>	$^6S_{5/2}$	7.4340
26	Fe	Iron	(Ar) 3d <sup>6</sup> 4s <sup>2</sup>	$^5D_4$	7.9024
27	Co	Cobalt	(Ar) 3d <sup>7</sup> 4s <sup>2</sup>	$^4F_{9/2}$	7.8810
28	Ni	Nickel	(Ar) 3d <sup>8</sup> 4s <sup>2</sup>	$^3F_4$	7.6399
29	Cu	Copper	(Ar) 3d <sup>10</sup> 4s	$^2S_{1/2}$	7.7264
30	Zn	Zinc	(Ar) 3d <sup>10</sup> 4s <sup>2</sup>	$^1S_0$	9.3942
31	Ga	Gallium	(Ar) 3d <sup>10</sup> 4s <sup>2</sup> 4p	$^2P_{1/2}$	5.9993
32	Ge	Germanium	(Ar) 3d <sup>10</sup> 4s <sup>2</sup> 4p <sup>2</sup>	$^3P_0$	7.8994
33	As	Arsenic	(Ar) 3d <sup>10</sup> 4s <sup>2</sup> 4p <sup>3</sup>	$^4S_{3/2}$	9.7886
34	Se	Selenium	(Ar) 3d <sup>10</sup> 4s <sup>2</sup> 4p <sup>4</sup>	$^3P_2$	9.7524
35	Br	Bromine	(Ar) 3d <sup>10</sup> 4s <sup>2</sup> 4p <sup>5</sup>	$^2P_{3/2}$	11.8138
36	Kr	Krypton	(Ar) 3d <sup>10</sup> 4s <sup>2</sup> 4p <sup>6</sup>	$^1S_0$	13.9996
37	Rb	Rubidium	(Kr) 5s	$^2S_{1/2}$	4.1771
38	Sr	Strontium	(Kr) 5s <sup>2</sup>	$^1S_0$	5.6949
39	Y	Yttrium	(Kr) 4d 5s <sup>2</sup>	$^2D_{3/2}$	6.2173
40	Zr	Zirconium	(Kr) 4d <sup>2</sup> 5s <sup>2</sup>	$^3F_2$	6.6339
41	Nb	Niobium	(Kr) 4d <sup>4</sup> 5s	$^6D_{1/2}$	6.7589
42	Mo	Molybdenum	(Kr) 4d <sup>5</sup> 5s	$^7S_3$	7.0924
43	Tc	Technetium	(Kr) 4d <sup>5</sup> 5s <sup>2</sup>	$^6S_{5/2}$	7.28
44	Ru	Ruthenium	(Kr) 4d <sup>7</sup> 5s	$^5F_5$	7.3605
45	Rh	Rhodium	(Kr) 4d <sup>8</sup> 5s	$^4F_{9/2}$	7.4589
46	Pd	Palladium	(Kr) 4d <sup>10</sup>	$^1S_0$	8.3369
47	Ag	Silver	(Kr) 4d <sup>10</sup> 5s	$^2S_{1/2}$	7.5762
48	Cd	Cadmium	(Kr) 4d <sup>10</sup> 5s <sup>2</sup>	$^1S_0$	8.9938

49	In	Indium	(Kr)4d <sup>10</sup> 5s <sup>2</sup>	5p		<sup>2</sup> P <sub>1/2</sub>	5.7864
50	Sn	Tin	(Kr)4d <sup>10</sup> 5s <sup>2</sup>	5p <sup>2</sup>		<sup>3</sup> P <sub>0</sub>	7.3439
51	Sb	Antimony	(Kr)4d <sup>10</sup> 5s <sup>2</sup>	5p <sup>3</sup>		<sup>4</sup> S <sub>3/2</sub>	8.6084
52	Te	Tellurium	(Kr)4d <sup>10</sup> 5s <sup>2</sup>	5p <sup>4</sup>		<sup>3</sup> P <sub>2</sub>	9.0096
53	I	Iodine	(Kr)4d <sup>10</sup> 5s <sup>2</sup>	5p <sup>5</sup>		<sup>2</sup> P <sub>3/2</sub>	10.4513
54	Xe	Xenon	(Kr)4d <sup>10</sup> 5s <sup>2</sup>	5p <sup>6</sup>		<sup>1</sup> S <sub>0</sub>	12.1298
55	Cs	Cesium	(Xe)	6s		<sup>2</sup> S <sub>1/2</sub>	3.8939
56	Ba	Barium	(Xe)	6s <sup>2</sup>		<sup>1</sup> S <sub>0</sub>	5.2117
57	La	Lanthanum	(Xe)	5d	6s <sup>2</sup>	<sup>2</sup> D <sub>3/2</sub>	5.5769
58	Ce	Cerium	(Xe)4f	5d	6s <sup>2</sup>	<sup>1</sup> G <sub>4</sub>	5.5387
59	Pr	Praseodymium	(Xe)4f <sup>3</sup>		6s <sup>2</sup>	<sup>4</sup> I <sub>9/2</sub>	5.473
60	Nd	Neodymium	(Xe)4f <sup>4</sup>		6s <sup>2</sup>	<sup>5</sup> I <sub>4</sub>	5.5250
61	Pm	Promethium	(Xe)4f <sup>5</sup>		6s <sup>2</sup>	<sup>6</sup> H <sub>5/2</sub>	5.582
62	Sm	Samarium	(Xe)4f <sup>6</sup>		6s <sup>2</sup>	<sup>7</sup> F <sub>0</sub>	5.6437
63	Eu	Europium	(Xe)4f <sup>7</sup>		6s <sup>2</sup>	<sup>8</sup> S <sub>7/2</sub>	5.6704
64	Gd	Gadolinium	(Xe)4f <sup>7</sup>	5d	6s <sup>2</sup>	<sup>9</sup> D <sub>2</sub>	6.1498
65	Tb	Terbium	(Xe)4f <sup>9</sup>		6s <sup>2</sup>	<sup>6</sup> H <sub>15/2</sub>	5.8638
66	Dy	Dysprosium	(Xe)4f <sup>10</sup>		6s <sup>2</sup>	<sup>5</sup> I <sub>8</sub>	5.9389
67	Ho	Holmium	(Xe)4f <sup>11</sup>		6s <sup>2</sup>	<sup>4</sup> I <sub>15/2</sub>	6.0215
68	Er	Erbium	(Xe)4f <sup>12</sup>		6s <sup>2</sup>	<sup>3</sup> H <sub>6</sub>	6.1077
69	Tm	Thulium	(Xe)4f <sup>13</sup>		6s <sup>2</sup>	<sup>2</sup> F <sub>7/2</sub>	6.1843
70	Yb	Ytterbium	(Xe)4f <sup>14</sup>		6s <sup>2</sup>	<sup>1</sup> S <sub>0</sub>	6.2542
71	Lu	Lutetium	(Xe)4f <sup>14</sup>	5d	6s <sup>2</sup>	<sup>2</sup> D <sub>3/2</sub>	5.4259
72	Hf	Hafnium	(Xe)4f <sup>14</sup>	5d <sup>2</sup>	6s <sup>2</sup>	<sup>3</sup> F <sub>2</sub>	6.8251
73	Ta	Tantalum	(Xe)4f <sup>14</sup>	5d <sup>3</sup>	6s <sup>2</sup>	<sup>4</sup> F <sub>3/2</sub>	7.5496
74	W	Tungsten	(Xe)4f <sup>14</sup>	5d <sup>4</sup>	6s <sup>2</sup>	<sup>5</sup> D <sub>0</sub>	7.8640
75	Re	Rhenium	(Xe)4f <sup>14</sup>	5d <sup>5</sup>	6s <sup>2</sup>	<sup>6</sup> S <sub>5/2</sub>	7.8335
76	Os	Osmium	(Xe)4f <sup>14</sup>	5d <sup>6</sup>	6s <sup>2</sup>	<sup>5</sup> D <sub>4</sub>	8.4382
77	Ir	Iridium	(Xe)4f <sup>14</sup>	5d <sup>7</sup>	6s <sup>2</sup>	<sup>4</sup> F <sub>9/2</sub>	8.9670
78	Pt	Platinum	(Xe)4f <sup>14</sup>	5d <sup>9</sup>	6s	<sup>3</sup> D <sub>3</sub>	8.9588
79	Au	Gold	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s	<sup>2</sup> S <sub>1/2</sub>	9.2255
80	Hg	Mercury	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup>	<sup>1</sup> S <sub>0</sub>	10.4375
81	Tl	Thallium	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup> 6p	<sup>2</sup> P <sub>1/2</sub>	6.1082
82	Pb	Lead	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup> 6p <sup>2</sup>	<sup>3</sup> P <sub>0</sub>	7.4167
83	Bi	Bismuth	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup> 6p <sup>3</sup>	<sup>4</sup> S <sub>3/2</sub>	7.2855
84	Po	Polonium	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup> 6p <sup>4</sup>	<sup>3</sup> P <sub>2</sub>	8.414
85	At	Astatine	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup> 6p <sup>5</sup>	<sup>2</sup> P <sub>3/2</sub>	
86	Rn	Radon	(Xe)4f <sup>14</sup>	5d <sup>10</sup>	6s <sup>2</sup> 6p <sup>6</sup>	<sup>1</sup> S <sub>0</sub>	10.7485
87	Fr	Francium	(Rn)		7s	<sup>2</sup> S <sub>1/2</sub>	4.0727
88	Ra	Radium	(Rn)		7s <sup>2</sup>	<sup>1</sup> S <sub>0</sub>	5.2784
89	Ac	Actinium	(Rn)	6d	7s <sup>2</sup>	<sup>2</sup> D <sub>3/2</sub>	5.3807
90	Th	Thorium	(Rn)	6d <sup>2</sup>	7s <sup>2</sup>	<sup>3</sup> F <sub>2</sub>	6.3067
91	Pa	Protactinium	(Rn)5f <sup>2</sup>	6d	7s <sup>2</sup>	<sup>4</sup> K <sub>11/2</sub> *	5.89
92	U	Uranium	(Rn)5f <sup>3</sup>	6d	7s <sup>2</sup>	<sup>5</sup> L <sub>6</sub> *	6.1939
93	Np	Neptunium	(Rn)5f <sup>4</sup>	6d	7s <sup>2</sup>	<sup>6</sup> L <sub>11/2</sub> *	6.2657
94	Pu	Plutonium	(Rn)5f <sup>6</sup>		7s <sup>2</sup>	<sup>7</sup> F <sub>0</sub>	6.0260
95	Am	Americium	(Rn)5f <sup>7</sup>		7s <sup>2</sup>	<sup>8</sup> S <sub>7/2</sub>	5.9738
96	Cm	Curium	(Rn)5f <sup>7</sup>	6d	7s <sup>2</sup>	<sup>9</sup> D <sub>2</sub>	5.9914
97	Bk	Berkelium	(Rn)5f <sup>9</sup>		7s <sup>2</sup>	<sup>6</sup> H <sub>15/2</sub>	6.1979
98	Cf	Californium	(Rn)5f <sup>10</sup>		7s <sup>2</sup>	<sup>5</sup> I <sub>8</sub>	6.2817
99	Es	Einsteinium	(Rn)5f <sup>11</sup>		7s <sup>2</sup>	<sup>4</sup> I <sub>15/2</sub>	6.3676
100	Fm	Fermium	(Rn)5f <sup>12</sup>		7s <sup>2</sup>	<sup>3</sup> H <sub>6</sub>	6.50
101	Md	Mendelevium	(Rn)5f <sup>13</sup>		7s <sup>2</sup>	<sup>2</sup> F <sub>7/2</sub>	6.58
102	No	Nobelium	(Rn)5f <sup>14</sup>		7s <sup>2</sup>	<sup>1</sup> S <sub>0</sub>	6.65
103	Lr	Lawrencium	(Rn)5f <sup>14</sup>		7s <sup>2</sup> 7p?	<sup>2</sup> P <sub>1/2</sub> ?	4.9?
104	Rf	Rutherfordium	(Rn)5f <sup>14</sup>	6d <sup>2</sup>	7s <sup>2</sup> ?	<sup>3</sup> F <sub>2</sub> ?	6.0?

\* The usual *LS* coupling scheme does not apply for these three elements. See the introductory note to the NIST table from which this table is taken.

## 6. ATOMIC AND NUCLEAR PROPERTIES OF MATERIALS

**Table 6.1** Abridged from [pdg.lbl.gov/AtomicNuclearProperties](http://pdg.lbl.gov/AtomicNuclearProperties) by D.E. Groom (2015). See web pages for more detail about entries in this table and for several hundred other substances. Parentheses in the  $dE/dx$  and density columns indicate gases at 20° C and 1 atm. Boiling points are at 1 atm. Refractive indices  $n$  are evaluated at the sodium D line blend (589.2 nm); values  $\gg 1$  in brackets indicate  $(n - 1) \times 10^6$  for gases at 0° C and 1 atm.

Material	$Z$	$A$	$\langle Z/A \rangle$	Nucl.coll. length $\lambda_T$ {g cm <sup>-2</sup> }	Nucl.inter. length $\lambda_I$ {g cm <sup>-2</sup> }	Rad.len. $X_0$ {g cm <sup>-2</sup> }	$dE/dx _{\min}$ { MeV g <sup>-1</sup> cm <sup>2</sup> }	Density {g cm <sup>-3</sup> ({gℓ <sup>-1</sup> })}	Melting point (K)	Boiling point (K)	Refract. index @ Na D
H <sub>2</sub>	1	1.008(7)	0.99212	42.8	52.0	63.04	(4.103)	0.071(0.084)	13.81	20.28	1.11[132.]
D <sub>2</sub>	1	2.01410177803(8)	0.49650	51.3	71.8	125.97	(2.053)	0.169(0.168)	18.7	23.65	1.11[138.]
He	2	4.002602(2)	0.49967	51.8	71.0	94.32	(1.937)	0.125(0.166)		4.220	1.02[35.0]
Li	3	6.94(2)	0.43221	52.2	71.3	82.78	1.639	0.534	453.6	1615.	
Be	4	9.0121831(5)	0.44384	55.3	77.8	65.19	1.595	1.848	1560.	2744.	
C diamond	6	12.0107(8)	0.49955	59.2	85.8	42.70	1.725	3.520			2.42
C graphite	6	12.0107(8)	0.49955	59.2	85.8	42.70	1.742	2.210			
N <sub>2</sub>	7	14.007(2)	0.49976	61.1	89.7	37.99	(1.825)	0.807(1.165)	63.15	77.29	1.20[298.]
O <sub>2</sub>	8	15.999(3)	0.50002	61.3	90.2	34.24	(1.801)	1.141(1.332)	54.36	90.20	1.22[271.]
F <sub>2</sub>	9	18.998403163(6)	0.47372	65.0	97.4	32.93	(1.676)	1.507(1.580)	53.53	85.03	[195.]
Ne	10	20.1797(6)	0.49555	65.7	99.0	28.93	(1.724)	1.204(0.839)	24.56	27.07	1.09[67.1]
Al	13	26.9815385(7)	0.48181	69.7	107.2	24.01	1.615	2.699	933.5	2792.	
Si	14	28.0855(3)	0.49848	70.2	108.4	21.82	1.664	2.329	1687.	3538.	3.95
Cl <sub>2</sub>	17	35.453(2)	0.47951	73.8	115.7	19.28	(1.630)	1.574(2.980)	171.6	239.1	[773.]
Ar	18	39.948(1)	0.45059	75.7	119.7	19.55	(1.519)	1.396(1.662)	83.81	87.26	1.23[281.]
Ti	22	47.867(1)	0.45961	78.8	126.2	16.16	1.477	4.540	1941.	3560.	
Fe	26	55.845(2)	0.46557	81.7	132.1	13.84	1.451	7.874	1811.	3134.	
Cu	29	63.546(3)	0.45636	84.2	137.3	12.86	1.403	8.960	1358.	2835.	
Ge	32	72.630(1)	0.44053	86.9	143.0	12.25	1.370	5.323	1211.	3106.	
Sn	50	118.710(7)	0.42119	98.2	166.7	8.82	1.263	7.310	505.1	2875.	
Xe	54	131.293(6)	0.41129	100.8	172.1	8.48	(1.255)	2.953(5.483)	161.4	165.1	1.39[701.]
W	74	183.84(1)	0.40252	110.4	191.9	6.76	1.145	19.300	3695.	5828.	
Pt	78	195.084(9)	0.39983	112.2	195.7	6.54	1.128	21.450	2042.	4098.	
Au	79	196.966569(5)	0.40108	112.5	196.3	6.46	1.134	19.320	1337.	3129.	
Pb	82	207.2(1)	0.39575	114.1	199.6	6.37	1.122	11.350	600.6	2022.	
U	92	[238.02891(3)]	0.38651	118.6	209.0	6.00	1.081	18.950	1408.	4404.	
Air (dry, 1 atm)			0.49919	61.3	90.1	36.62	(1.815)	(1.205)		78.80	[289]
Shielding concrete			0.50274	65.1	97.5	26.57	1.711	2.300			
Borosilicate glass (Pyrex)			0.49707	64.6	96.5	28.17	1.696	2.230			
Lead glass			0.42101	95.9	158.0	7.87	1.255	6.220			
Standard rock			0.50000	66.8	101.3	26.54	1.688	2.650			
Methane (CH <sub>4</sub> )			0.62334	54.0	73.8	46.47	(2.417)	(0.667)	90.68	111.7	[444.]
Ethane (C <sub>2</sub> H <sub>6</sub> )			0.59861	55.0	75.9	45.66	(2.304)	(1.263)	90.36	184.5	
Propane (C <sub>3</sub> H <sub>8</sub> )			0.58962	55.3	76.7	45.37	(2.262)	0.493(1.868)	85.52	231.0	
Butane (C <sub>4</sub> H <sub>10</sub> )			0.59497	55.5	77.1	45.23	(2.278)	(2.489)	134.9	272.6	
Octane (C <sub>8</sub> H <sub>18</sub> )			0.57778	55.8	77.8	45.00	2.123	0.703	214.4	398.8	
Paraffin (CH <sub>3</sub> (CH <sub>2</sub> ) <sub>n≈23</sub> CH <sub>3</sub> )			0.57275	56.0	78.3	44.85	2.088	0.930			
Nylon (type 6, 6/6)			0.54790	57.5	81.6	41.92	1.973	1.18			
Polycarbonate (Lexan)			0.52697	58.3	83.6	41.50	1.886	1.20			
Polyethylene ([CH <sub>2</sub> CH <sub>2</sub> ] <sub>n</sub> )			0.57034	56.1	78.5	44.77	2.079	0.89			
Polyethylene terephthalate (Mylar)			0.52037	58.9	84.9	39.95	1.848	1.40			
Polyimide film (Kapton)			0.51264	59.2	85.5	40.58	1.820	1.42			
Polymethylmethacrylate (acrylic)			0.53937	58.1	82.8	40.55	1.929	1.19			1.49
Polypropylene			0.55998	56.1	78.5	44.77	2.041	0.90			
Polystyrene ([C <sub>6</sub> H <sub>5</sub> CHCH <sub>2</sub> ] <sub>n</sub> )			0.53768	57.5	81.7	43.79	1.936	1.06			1.59
Polytetrafluoroethylene (Teflon)			0.47992	63.5	94.4	34.84	1.671	2.20			
Polyvinyltoluene			0.54141	57.3	81.3	43.90	1.956	1.03			1.58
Aluminum oxide (sapphire)			0.49038	65.5	98.4	27.94	1.647	3.970	2327.	3273.	1.77
Barium fluoride (BaF <sub>2</sub> )			0.42207	90.8	149.0	9.91	1.303	4.893	1641.	2533.	1.47
Bismuth germanate (BGO)			0.42065	96.2	159.1	7.97	1.251	7.130	1317.		2.15
Carbon dioxide gas (CO <sub>2</sub> )			0.49989	60.7	88.9	36.20	1.819	(1.842)			[449.]
Solid carbon dioxide (dry ice)			0.49989	60.7	88.9	36.20	1.787	1.563	Sublimes at 194.7 K		
Cesium iodide (CsI)			0.41569	100.6	171.5	8.39	1.243	4.510	894.2	1553.	1.79
Lithium fluoride (LiF)			0.46262	61.0	88.7	39.26	1.614	2.635	1121.	1946.	1.39
Lithium hydride (LiH)			0.50321	50.8	68.1	79.62	1.897	0.820	965.		
Lead tungstate (PbWO <sub>4</sub> )			0.41315	100.6	168.3	7.39	1.229	8.300	1403.		2.20
Silicon dioxide (SiO <sub>2</sub> , fused quartz)			0.49930	65.2	97.8	27.05	1.699	2.200	1986.	3223.	1.46
Sodium chloride (NaCl)			0.47910	71.2	110.1	21.91	1.847	2.170	1075.	1738.	1.54
Sodium iodide (NaI)			0.42697	93.1	154.6	9.49	1.305	3.667	933.2	1577.	1.77
Water (H <sub>2</sub> O)			0.55509	58.5	83.3	36.08	1.992	1.000	273.1	373.1	1.33
Silica aerogel			0.50093	65.0	97.3	27.25	1.740	0.200	(0.03 H <sub>2</sub> O, 0.97 SiO <sub>2</sub> )		

Material	Dielectric constant ( $\kappa = \epsilon/\epsilon_0$ ) ( ) is $(\kappa-1)\times 10^6$ for gas	Young's modulus [ $10^6$ psi]	Coeff. of thermal expansion [ $10^{-6}\text{cm/cm-}^\circ\text{C}$ ]	Specific heat [cal/g- $^\circ\text{C}$ ]	Electrical resistivity [ $\mu\Omega\text{cm}(@^\circ\text{C})$ ]	Thermal conductivity [cal/cm- $^\circ\text{C-sec}$ ]
H <sub>2</sub>	(253.9)	—	—	—	—	—
He	(64)	—	—	—	—	—
Li	—	—	56	0.86	8.55(0 $^\circ$ )	0.17
Be	—	37	12.4	0.436	5.885(0 $^\circ$ )	0.38
C	—	0.7	0.6–4.3	0.165	1375(0 $^\circ$ )	0.057
N <sub>2</sub>	(548.5)	—	—	—	—	—
O <sub>2</sub>	(495)	—	—	—	—	—
Ne	(127)	—	—	—	—	—
Al	—	10	23.9	0.215	2.65(20 $^\circ$ )	0.53
Si	11.9	16	2.8–7.3	0.162	—	0.20
Ar	(517)	—	—	—	—	—
Ti	—	16.8	8.5	0.126	50(0 $^\circ$ )	—
Fe	—	28.5	11.7	0.11	9.71(20 $^\circ$ )	0.18
Cu	—	16	16.5	0.092	1.67(20 $^\circ$ )	0.94
Ge	16.0	—	5.75	0.073	—	0.14
Sn	—	6	20	0.052	11.5(20 $^\circ$ )	0.16
Xe	—	—	—	—	—	—
W	—	50	4.4	0.032	5.5(20 $^\circ$ )	0.48
Pt	—	21	8.9	0.032	9.83(0 $^\circ$ )	0.17
Pb	—	2.6	29.3	0.038	20.65(20 $^\circ$ )	0.083
U	—	—	36.1	0.028	29(20 $^\circ$ )	0.064

## 7. ELECTROMAGNETIC RELATIONS

Revised September 2005 by H.G. Spieler (LBNL).

Quantity	Gaussian CGS	SI
Conversion factors: Charge: Potential: Magnetic field:	$2.997\,924\,58 \times 10^9$ esu $(1/299.792\,458)$ statvolt (ergs/esu) $10^4$ gauss = $10^4$ dyne/esu	$= 1\text{ C} = 1\text{ A s}$ $= 1\text{ V} = 1\text{ J C}^{-1}$ $= 1\text{ T} = 1\text{ N A}^{-1}\text{m}^{-1}$
	$\mathbf{F} = q(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B})$	$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$
	$\nabla \cdot \mathbf{D} = 4\pi\rho$ $\nabla \times \mathbf{H} - \frac{1}{c} \frac{\partial \mathbf{D}}{\partial t} = \frac{4\pi}{c} \mathbf{J}$ $\nabla \cdot \mathbf{B} = 0$ $\nabla \times \mathbf{E} + \frac{1}{c} \frac{\partial \mathbf{B}}{\partial t} = 0$	$\nabla \cdot \mathbf{D} = \rho$ $\nabla \times \mathbf{H} - \frac{\partial \mathbf{D}}{\partial t} = \mathbf{J}$ $\nabla \cdot \mathbf{B} = 0$ $\nabla \times \mathbf{E} + \frac{\partial \mathbf{B}}{\partial t} = 0$
Constitutive relations:	$\mathbf{D} = \mathbf{E} + 4\pi\mathbf{P}, \quad \mathbf{H} = \mathbf{B} - 4\pi\mathbf{M}$	$\mathbf{D} = \epsilon_0\mathbf{E} + \mathbf{P}, \quad \mathbf{H} = \mathbf{B}/\mu_0 - \mathbf{M}$
Linear media:	$\mathbf{D} = \epsilon\mathbf{E}, \quad \mathbf{H} = \mathbf{B}/\mu$ 1 1	$\mathbf{D} = \epsilon\mathbf{E}, \quad \mathbf{H} = \mathbf{B}/\mu$ $\epsilon_0 = 8.854\,187 \dots \times 10^{-12} \text{ F m}^{-1}$ $\mu_0 = 4\pi \times 10^{-7} \text{ N A}^{-2}$
	$\mathbf{E} = -\nabla V - \frac{1}{c} \frac{\partial \mathbf{A}}{\partial t}$ $\mathbf{B} = \nabla \times \mathbf{A}$	$\mathbf{E} = -\nabla V - \frac{\partial \mathbf{A}}{\partial t}$ $\mathbf{B} = \nabla \times \mathbf{A}$
	$V = \sum_{\text{charges}} \frac{q_i}{r_i} = \int \frac{\rho(\mathbf{r}')}{ \mathbf{r} - \mathbf{r}' } d^3x'$ $\mathbf{A} = \frac{1}{c} \oint \frac{I d\boldsymbol{\ell}}{ \mathbf{r} - \mathbf{r}' } = \frac{1}{c} \int \frac{\mathbf{J}(\mathbf{r}')}{ \mathbf{r} - \mathbf{r}' } d^3x'$	$V = \frac{1}{4\pi\epsilon_0} \sum_{\text{charges}} \frac{q_i}{r_i} = \frac{1}{4\pi\epsilon_0} \int \frac{\rho(\mathbf{r}')}{ \mathbf{r} - \mathbf{r}' } d^3x'$ $\mathbf{A} = \frac{\mu_0}{4\pi} \oint \frac{I d\boldsymbol{\ell}}{ \mathbf{r} - \mathbf{r}' } = \frac{\mu_0}{4\pi} \int \frac{\mathbf{J}(\mathbf{r}')}{ \mathbf{r} - \mathbf{r}' } d^3x'$
	$\mathbf{E}'_{\parallel} = \mathbf{E}_{\parallel}$ $\mathbf{E}'_{\perp} = \gamma(\mathbf{E}_{\perp} + \frac{1}{c} \mathbf{v} \times \mathbf{B})$ $\mathbf{B}'_{\parallel} = \mathbf{B}_{\parallel}$ $\mathbf{B}'_{\perp} = \gamma(\mathbf{B}_{\perp} - \frac{1}{c} \mathbf{v} \times \mathbf{E})$	$\mathbf{E}'_{\parallel} = \mathbf{E}_{\parallel}$ $\mathbf{E}'_{\perp} = \gamma(\mathbf{E}_{\perp} + \mathbf{v} \times \mathbf{B})$ $\mathbf{B}'_{\parallel} = \mathbf{B}_{\parallel}$ $\mathbf{B}'_{\perp} = \gamma(\mathbf{B}_{\perp} - \frac{1}{c^2} \mathbf{v} \times \mathbf{E})$
$\frac{1}{4\pi\epsilon_0} = c^2 \times 10^{-7} \text{ N A}^{-2} = 8.987\,55 \dots \times 10^9 \text{ m F}^{-1}; \quad \frac{\mu_0}{4\pi} = 10^{-7} \text{ N A}^{-2}; \quad c = \frac{1}{\sqrt{\mu_0\epsilon_0}} = 2.997\,924\,58 \times 10^8 \text{ m s}^{-1}$		



### 7.1. Impedances (SI units)

$\rho$  = resistivity at room temperature in  $10^{-8} \Omega \text{ m}$ :

$\sim 1.7$ for Cu	$\sim 5.5$ for W
$\sim 2.4$ for Au	$\sim 73$ for SS 304
$\sim 2.8$ for Al	$\sim 100$ for Nichrome

(Al alloys may have double the Al value.)

For alternating currents, instantaneous current  $I$ , voltage  $V$ , angular frequency  $\omega$ :

$$V = V_0 e^{j\omega t} = ZI. \quad (7.1)$$

Impedance of self-inductance  $L$ :  $Z = j\omega L$ .

Impedance of capacitance  $C$ :  $Z = 1/j\omega C$ .

Impedance of free space:  $Z = \sqrt{\mu_0/\epsilon_0} = 376.7 \Omega$ .

High-frequency surface impedance of a good conductor:

$$Z = \frac{(1+j)\rho}{\delta}, \quad \text{where } \delta = \text{skin depth}; \quad (7.2)$$

$$\delta = \sqrt{\frac{\rho}{\pi\nu\mu}} \approx \frac{6.6 \text{ cm}}{\sqrt{\nu \text{ (Hz)}}} \quad \text{for Cu}. \quad (7.3)$$

### 7.2. Capacitors, inductors, and transmission Lines

The capacitance between two parallel plates of area  $A$  spaced by the distance  $d$  and enclosing a medium with the dielectric constant  $\epsilon$  is

$$C = K\epsilon A/d, \quad (7.4)$$

where the correction factor  $K$  depends on the extent of the fringing field. If the dielectric fills the capacitor volume without extending beyond the electrodes, the correction factor  $K \approx 0.8$  for capacitors of typical geometry.

The inductance at high frequencies of a straight wire whose length  $\ell$  is much greater than the wire diameter  $d$  is

$$L \approx 2.0 \left[ \frac{\text{nH}}{\text{cm}} \right] \cdot \ell \left( \ln \left( \frac{4\ell}{d} \right) - 1 \right). \quad (7.5)$$

For very short wires, representative of vias in a printed circuit board, the inductance is

$$L(\text{in nH}) \approx \ell/d. \quad (7.6)$$

A transmission line is a pair of conductors with inductance  $L$  and capacitance  $C$ . The characteristic impedance  $Z = \sqrt{L/C}$  and the phase velocity  $v_p = 1/\sqrt{LC} = 1/\sqrt{\mu\epsilon}$ , which decreases with the inverse square root of the dielectric constant of the medium. Typical coaxial and ribbon cables have a propagation delay of about 5 ns/cm. The impedance of a coaxial cable with outer diameter  $D$  and inner diameter  $d$  is

$$Z = 60 \Omega \cdot \frac{1}{\sqrt{\epsilon_r}} \ln \frac{D}{d}, \quad (7.7)$$

where the relative dielectric constant  $\epsilon_r = \epsilon/\epsilon_0$ . A pair of parallel wires of diameter  $d$  and spacing  $a > 2.5d$  has the impedance

$$Z = 120 \Omega \cdot \frac{1}{\sqrt{\epsilon_r}} \ln \frac{2a}{d}. \quad (7.8)$$

This yields the impedance of a wire at a spacing  $h$  above a ground plane,

$$Z = 60 \Omega \cdot \frac{1}{\sqrt{\epsilon_r}} \ln \frac{4h}{d}. \quad (7.9)$$

A common configuration utilizes a thin rectangular conductor above a ground plane with an intermediate dielectric (microstrip). Detailed calculations for this and other transmission line configurations are given by Gunston.\*

### 7.3. Synchrotron radiation (CGS units)

For a particle of charge  $e$ , velocity  $v = \beta c$ , and energy  $E = \gamma mc^2$ , traveling in a circular orbit of radius  $R$ , the classical energy loss per revolution  $\delta E$  is

$$\delta E = \frac{4\pi}{3} \frac{e^2}{R} \beta^3 \gamma^4. \quad (7.10)$$

For high-energy electrons or positrons ( $\beta \approx 1$ ), this becomes

$$\delta E \text{ (in MeV)} \approx 0.0885 [E(\text{in GeV})]^4 / R(\text{in m}). \quad (7.11)$$

For  $\gamma \gg 1$ , the energy radiated per revolution into the photon energy interval  $d(\hbar\omega)$  is

$$dI = \frac{8\pi}{9} \alpha \gamma F(\omega/\omega_c) d(\hbar\omega), \quad (7.12)$$

where  $\alpha = e^2/\hbar c$  is the fine-structure constant and

$$\omega_c = \frac{3\gamma^3 c}{2R} \quad (7.13)$$

is the critical frequency. The normalized function  $F(y)$  is

$$F(y) = \frac{9}{8\pi} \sqrt{3} y \int_y^\infty K_{5/3}(x) dx, \quad (7.14)$$

where  $K_{5/3}(x)$  is a modified Bessel function of the third kind. For electrons or positrons,

$$\hbar\omega_c \text{ (in keV)} \approx 2.22 [E(\text{in GeV})]^3 / R(\text{in m}). \quad (7.15)$$

Fig. 7.1 shows  $F(y)$  over the important range of  $y$ .

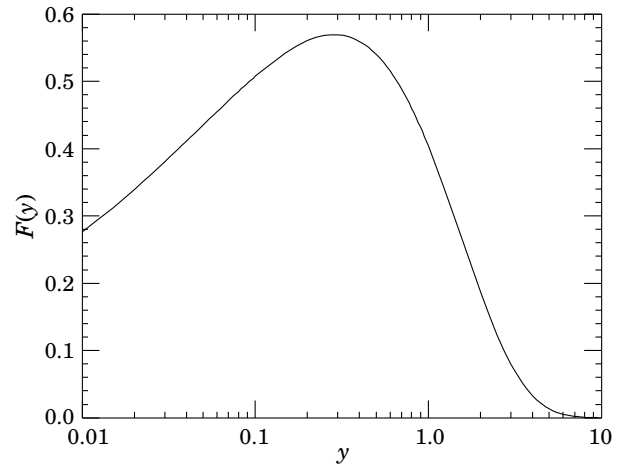


Figure 7.1: The normalized synchrotron radiation spectrum  $F(y)$ .

For  $\gamma \gg 1$  and  $\omega \ll \omega_c$ ,

$$\frac{dI}{d(\hbar\omega)} \approx 3.3\alpha (\omega R/c)^{1/3}, \quad (7.16)$$

whereas for

$$\gamma \gg 1 \text{ and } \omega \gtrsim 3\omega_c,$$

$$\frac{dI}{d(\hbar\omega)} \approx \sqrt{\frac{3\pi}{2}} \alpha \gamma \left( \frac{\omega}{\omega_c} \right)^{1/2} e^{-\omega/\omega_c} \left[ 1 + \frac{55}{72} \frac{\omega_c}{\omega} + \dots \right]. \quad (7.17)$$

The radiation is confined to angles  $\lesssim 1/\gamma$  relative to the instantaneous direction of motion. For  $\gamma \gg 1$ , where Eq. (7.12) applies, the mean number of photons emitted per revolution is

$$N_\gamma = \frac{5\pi}{\sqrt{3}} \alpha \gamma, \quad (7.18)$$

and the mean energy per photon is

$$\langle \hbar\omega \rangle = \frac{8}{15\sqrt{3}} \hbar\omega_c. \quad (7.19)$$

When  $\langle \hbar\omega \rangle \gtrsim O(E)$ , quantum corrections are important.

\* M.A.R. Gunston. Microwave Transmission Line Data, Noble Publishing Corp., Atlanta (1997) ISBN 1-884932-57-6, TK6565.T73G85.

See J.D. Jackson, *Classical Electrodynamics*, 3rd edition (John Wiley & Sons, New York, 1998) for more formulae and details. (Note that earlier editions had  $\omega_c$  twice as large as Eq. (7.13).)

## 8. NAMING SCHEME FOR HADRONS

Revised 2004 by M. Roos (University of Finland) and C.G. Wohl (LBNL).

## 8.1. Introduction

We introduced in the 1986 edition [1] a new naming scheme for the hadrons. Changes from older terminology affected mainly the heavier mesons made of the light ( $u$ ,  $d$ , and  $s$ ) quarks. Old and new names were listed alongside until 1994. Names also change from edition to edition because some characteristic like mass or spin changes. The Summary Tables give both the new and old names whenever a change occurred.

8.2. “Neutral-flavor” mesons ( $S=C=B=T=0$ )

Table 8.1 shows the names for mesons having the strangeness and all heavy-flavor quantum numbers equal to zero. The scheme is designed for all ordinary non-exotic mesons, but it will work for many exotic types too, if needed.

**Table 8.1:** Symbols for mesons with the strangeness and all heavy-flavor quantum numbers equal to zero.

$J^{PC}$		$0^{-+}$	$1^{+-}$	$1^{--}$	$0^{++}$
		$2^{-+}$	$3^{+-}$	$2^{--}$	$1^{++}$
		$\vdots$	$\vdots$	$\vdots$	$\vdots$
$q\bar{q}$ content	$^{2S+1}L_J = ^1(L\text{ even})_J \ ^1(L\text{ odd})_J \ ^3(L\text{ even})_J \ ^3(L\text{ odd})_J$				
$u\bar{d}, u\bar{u} - d\bar{d}, d\bar{u} \ (I=1)$	$\pi$	$b$	$\rho$	$a$	
$d\bar{d} + u\bar{u}$ and/or $s\bar{s}$ } $(I=0)$	$\eta, \eta'$	$h, h'$	$\omega, \phi$	$f, f'$	
$c\bar{c}$	$\eta_c$	$h_c$	$\psi^\dagger$	$\chi_c$	
$b\bar{b}$	$\eta_b$	$h_b$	$\Upsilon$	$\chi_b$	
$t\bar{t}$	$\eta_t$	$h_t$	$\theta$	$\chi_t$	

<sup>†</sup>The  $J/\psi$  remains the  $J/\psi$ .

First, we assign names to those states with quantum numbers compatible with being  $q\bar{q}$  states. The rows of the Table give the possible  $q\bar{q}$  content. The columns give the possible parity/charge-conjugation states,

$$PC = --, +-, --, \text{ and } ++;$$

these combinations correspond one-to-one with the angular-momentum state  $^{2S+1}L_J$  of the  $q\bar{q}$  system being

$$^1(L\text{ even})_J, ^1(L\text{ odd})_J, ^3(L\text{ even})_J, \text{ or } ^3(L\text{ odd})_J.$$

Here  $S$ ,  $L$ , and  $J$  are the spin, orbital, and total angular momenta of the  $q\bar{q}$  system. The quantum numbers are related by  $P = (-1)^{L+1}$ ,  $C = (-1)^{L+S}$ , and  $G$  parity  $= (-1)^{L+S+I}$ , where of course the  $C$  quantum number is only relevant to neutral mesons.

The entries in the Table give the meson names. The spin  $J$  is added as a subscript except for pseudoscalar and vector mesons, and the mass is added in parentheses for mesons that decay strongly. However, for the lightest meson resonances, we omit the mass.

Measurements of the mass, quark content (where relevant), and quantum numbers  $I$ ,  $J$ ,  $P$ , and  $C$  (or  $G$ ) of a meson thus fix its symbol. Conversely, these properties may be inferred unambiguously from the symbol.

If the main symbol cannot be assigned because the quantum numbers are unknown,  $X$  is used. Sometimes it is not known whether a meson is mainly the isospin-0 mix of  $u\bar{u}$  and  $d\bar{d}$  or is mainly  $s\bar{s}$ . A prime (or pair  $\omega$ ,  $\phi$ ) may be used to distinguish two such mixing states.

We follow custom and use spectroscopic names such as  $\Upsilon(1S)$  as the primary name for most of those  $\psi$ ,  $\Upsilon$ , and  $\chi$  states whose spectroscopic identity is known. We use the form  $\Upsilon(9460)$  as an alternative, and as the primary name when the spectroscopic identity is not known.

Names are assigned for  $t\bar{t}$  mesons, although the top quark is evidently so heavy that it is expected to decay too rapidly for bound states to form.

Gluonium states or other mesons that are not  $q\bar{q}$  states are, if the quantum numbers are *not* exotic, to be named just as are the  $q\bar{q}$  mesons. Such states will probably be difficult to distinguish from  $q\bar{q}$  states and will likely mix with them, and we make no attempt to distinguish those “mostly gluonium” from those “mostly  $q\bar{q}$ .”

An “exotic” meson with  $J^{PC}$  quantum numbers that a  $q\bar{q}$  system cannot have, namely  $J^{PC} = 0^{--}, 0^{+-}, 1^{-+}, 2^{+-}, 3^{-+}, \dots$ , would use the same symbol as does an ordinary meson with all the same quantum numbers as the exotic meson except for the  $C$  parity. But then the  $J$  subscript may still distinguish it; for example, an isospin-0  $1^{-+}$  meson could be denoted  $\omega_1$ .

8.3. Mesons with nonzero  $S$ ,  $C$ ,  $B$ , and/or  $T$ 

Since the strangeness or a heavy flavor of these mesons is nonzero, none of them are eigenstates of charge conjugation, and in each of them one of the quarks is heavier than the other. The rules are:

1. The main symbol is an upper-case italic letter indicating the heavier quark as follows:

$$s \rightarrow \bar{K} \quad c \rightarrow D \quad b \rightarrow \bar{B} \quad t \rightarrow T.$$

We use the convention that *the flavor and the charge of a quark have the same sign*. Thus the strangeness of the  $s$  quark is negative, the charm of the  $c$  quark is positive, and the bottom of the  $b$  quark is negative. In addition,  $I_3$  of the  $u$  and  $d$  quarks are positive and negative, respectively. The effect of this convention is as follows: *Any flavor carried by a charged meson has the same sign as its charge*. Thus the  $K^+$ ,  $D^+$ , and  $B^+$  have positive strangeness, charm, and bottom, respectively, and all have positive  $I_3$ . The  $D_s^+$  has positive charm *and* strangeness. Furthermore, the  $\Delta(\text{flavor}) = \Delta Q$  rule, best known for the kaons, applies to every flavor.

2. If the lighter quark is not a  $u$  or a  $d$  quark, its identity is given by a subscript. The  $D_s^+$  is an example.
3. If the spin-parity is in the “normal” series,  $J^P = 0^+, 1^-, 2^+, \dots$ , a superscript “\*” is added.
4. The spin is added as a subscript except for pseudoscalar or vector mesons.

## 8.4. Ordinary (3-quark) baryons

The symbols  $N$ ,  $\Delta$ ,  $\Lambda$ ,  $\Sigma$ ,  $\Xi$ , and  $\Omega$  used for more than 30 years for the baryons made of light quarks ( $u$ ,  $d$ , and  $s$  quarks) tell the isospin and quark content, and the same information is conveyed by the symbols used for the baryons containing one or more heavy quarks ( $c$  and  $b$  quarks). The rules are:

1. Baryons with *three*  $u$  and/or  $d$  quarks are  $N$ ’s (isospin 1/2) or  $\Delta$ ’s (isospin 3/2).
2. Baryons with *two*  $u$  and/or  $d$  quarks are  $\Lambda$ ’s (isospin 0) or  $\Sigma$ ’s (isospin 1). If the third quark is a  $c$ ,  $b$ , or  $t$  quark, its identity is given by a subscript.
3. Baryons with *one*  $u$  or  $d$  quark are  $\Xi$ ’s (isospin 1/2). One or two subscripts are used if one or both of the remaining quarks are heavy: thus  $\Xi_c$ ,  $\Xi_{cc}$ ,  $\Xi_b$ , *etc.*\*
4. Baryons with *no*  $u$  or  $d$  quarks are  $\Omega$ ’s (isospin 0), and subscripts indicate any heavy-quark content.
5. A baryon that decays strongly has its mass as part of its name. Thus  $p$ ,  $\Sigma^+$ ,  $\Omega^-$ ,  $\Lambda_c^+$ , *etc.*, but  $\Delta(1232)^0$ ,  $\Sigma(1385)^-$ ,  $\Xi_c(2645)^+$ , *etc.*

In short, the number of  $u$  plus  $d$  quarks together with the isospin determine the main symbol, and subscripts indicate any content of heavy quarks. A  $\Sigma$  always has isospin 1, an  $\Omega$  always has isospin 0, *etc.*

### 8.5. Exotic baryons

In 2003, several experiments reported finding a strangeness  $S = +1$ , charge  $Q = +1$  baryon, and one experiment reported finding an  $S = -2$ ,  $Q = -2$  baryon. Baryons with such quantum numbers cannot be made from three quarks, and thus they are exotic. The  $S = +1$  baryon, which once would have been called a  $Z$ , was quickly dubbed the  $\Theta(1540)^+$ , and we proposed to name the  $S = -2$  baryon the  $\Phi(1860)$ . However, these “discoveries” were then completely ruled out by many experiments with far larger statistics: See our 2008 *Review* [2].

#### Footnote and Reference:

\* Sometimes a prime is necessary to distinguish two  $\Xi_c$ 's in the same  $SU(n)$  multiplet. See the “Note on Charmed Baryons” in the Charmed Baryon Listings.

1. Particle Data Group: M. Aguilar-Benitez *et al.*, Phys. Lett. **170B** (1986).
2. Particle Data Group: C. Amsler *et al.*, Phys. Lett. **B667**, 1 (2008).

## 9. QUANTUM CHROMODYNAMICS

Revised September 2015 (April 2016 for section on  $\alpha_s$ ) by S. Bethke (Max-Planck-Institute of Physics, Munich), G. Dissertori (ETH Zurich), and G.P. Salam (CERN).<sup>1</sup>

### 9.1. Basics

Quantum Chromodynamics (QCD), the gauge field theory that describes the strong interactions of colored quarks and gluons, is the SU(3) component of the SU(3)×SU(2)×U(1) Standard Model of Particle Physics.

The Lagrangian of QCD is given by

$$\mathcal{L} = \sum_q \bar{\psi}_{q,a} (i\gamma^\mu \partial_\mu \delta_{ab} - g_s \gamma^\mu t_{ab}^C \mathcal{A}_\mu^C - m_q \delta_{ab}) \psi_{q,b} - \frac{1}{4} F_{\mu\nu}^A F^{A\mu\nu}, \quad (9.1)$$

where repeated indices are summed over. The  $\gamma^\mu$  are the Dirac  $\gamma$ -matrices. The  $\psi_{q,a}$  are quark-field spinors for a quark of flavor  $q$  and mass  $m_q$ , with a color-index  $a$  that runs from  $a = 1$  to  $N_c = 3$ , *i.e.* quarks come in three “colors.” Quarks are said to be in the fundamental representation of the SU(3) color group.

The  $\mathcal{A}_\mu^C$  correspond to the gluon fields, with  $C$  running from 1 to  $N_c^2 - 1 = 8$ , *i.e.* there are eight kinds of gluon. Gluons transform under the adjoint representation of the SU(3) color group. The  $t_{ab}^C$  correspond to eight  $3 \times 3$  matrices and are the generators of the SU(3) group (*cf.* the section on “SU(3) isoscalar factors and representation matrices” in this *Review*, with  $t_{ab}^C \equiv \lambda_{ab}^C/2$ ). They encode the fact that a gluon’s interaction with a quark rotates the quark’s color in SU(3) space. The quantity  $g_s$  is the QCD coupling constant. Finally, the field tensor  $F_{\mu\nu}^A$  is given by

$$F_{\mu\nu}^A = \partial_\mu \mathcal{A}_\nu^A - \partial_\nu \mathcal{A}_\mu^A - g_s f_{ABC} \mathcal{A}_\mu^B \mathcal{A}_\nu^C \quad [t^A, t^B] = i f_{ABC} t^C, \quad (9.2)$$

where the  $f_{ABC}$  are the structure constants of the SU(3) group.

Neither quarks nor gluons are observed as free particles. Hadrons are color-singlet (*i.e.* color-neutral) combinations of quarks, anti-quarks, and gluons.

Ab-initio predictive methods for QCD include lattice gauge theory and perturbative expansions in the coupling. The Feynman rules of QCD involve a quark-antiquark-gluon ( $q\bar{q}g$ ) vertex, a 3-gluon vertex (both proportional to  $g_s$ ), and a 4-gluon vertex (proportional to  $g_s^2$ ). A full set of Feynman rules is to be found for example in Ref. 1.

Useful color-algebra relations include:  $t_{ab}^A t_{bc}^A = C_F \delta_{ac}$ , where  $C_F \equiv (N_c^2 - 1)/(2N_c) = 4/3$  is the color-factor (“Casimir”) associated with gluon emission from a quark;  $f_{ACD} f_{BCD} = C_A \delta_{AB}$  where  $C_A \equiv N_c = 3$  is the color-factor associated with gluon emission from a gluon;  $t_{ab}^A t_{ab}^B = T_R \delta_{AB}$ , where  $T_R = 1/2$  is the color-factor for a gluon to split to a  $q\bar{q}$  pair.

The fundamental parameters of QCD are the coupling  $g_s$  (or  $\alpha_s = \frac{g_s^2}{4\pi}$ ) and the quark masses  $m_q$ .

There is freedom for an additional CP-violating term to be present in the QCD Lagrangian,  $\theta \frac{\alpha_s}{8\pi} F_{\mu\nu}^A \tilde{F}^{A\mu\nu}$ , where  $\tilde{F}^{A\mu\nu}$  is the dual of the gluon field tensor,  $\frac{1}{2} \epsilon_{\mu\nu\rho\sigma} F^{A\rho\sigma}$ , where  $\epsilon_{\mu\nu\rho\sigma}$  is the fully antisymmetric Levi-Cevita symbol. Experimental limits on the neutron electric dipole moment [2] constrain the coefficient of this contribution to satisfy  $|\theta| \lesssim 10^{-10}$ . Further discussion is to be found in Ref. 3 and in the Axions section in the Listings of this *Review*.

This section will concentrate mainly on perturbative aspects of QCD as they relate to collider physics. Related textbooks and reviews include Refs. 1,4–7. Aspects specific to Monte Carlo event generators are reviewed in the dedicated section 41. Lattice QCD is also reviewed in a section of its own, Sec. 18, with further discussion of perturbative and non-perturbative aspects to be found in the sections on “Quark Masses”, “The CKM quark-mixing matrix”, “Structure Functions”, “Fragmentation Functions”, and “Heavy-Quark and Soft-Collinear Effective Theory” in this *Review*. For an overview of some of the QCD issues and recent results in heavy-ion physics, see for example Refs. [8–10].

#### 9.1.1. Running coupling :

In the framework of perturbative QCD (pQCD), predictions for observables are expressed in terms of the renormalized coupling  $\alpha_s(\mu_R^2)$ , a function of an (unphysical) renormalization scale  $\mu_R$ . When one takes  $\mu_R$  close to the scale of the momentum transfer  $Q$  in a given process, then  $\alpha_s(\mu_R^2 \simeq Q^2)$  is indicative of the effective strength of the strong interaction in that process.

The coupling satisfies the following renormalization group equation (RGE):

$$\mu_R^2 \frac{d\alpha_s}{d\mu_R^2} = \beta(\alpha_s) = -(b_0 \alpha_s^2 + b_1 \alpha_s^3 + b_2 \alpha_s^4 + \dots) \quad (9.3)$$

where  $b_0 = (11C_A - 4n_f T_R)/(12\pi) = (33 - 2n_f)/(12\pi)$  is referred to as the 1-loop  $\beta$ -function coefficient, the 2-loop coefficient is  $b_1 = (17C_A^2 - n_f T_R(10C_A + 6C_F))/(24\pi^2) = (153 - 19n_f)/(24\pi^2)$ , and the 3-loop coefficient is  $b_2 = (2857 - \frac{5033}{9}n_f + \frac{325}{27}n_f^2)/(128\pi^3)$  for the SU(3) values of  $C_A$  and  $C_F$ . The 4-loop coefficient,  $b_3$ , is to be found in Refs. 11, 12. The coefficients  $b_2$  and  $b_3$  (and beyond) are renormalization-scheme-dependent, and given here in the modified minimal subtraction ( $\overline{\text{MS}}$ ) scheme [13], by far the most widely used scheme in QCD.

The minus sign in Eq. (9.3) is the origin of Asymptotic Freedom [14,15], *i.e.* the fact that the strong coupling becomes weak for processes involving large momentum transfers (“hard processes”). For momentum transfers in the 100 GeV – TeV range,  $\alpha_s \sim 0.1$ , while the theory is strongly interacting for scales around and below 1 GeV.

The  $\beta$ -function coefficients, the  $b_i$ , are given for the coupling of an *effective theory* in which  $n_f$  of the quark flavors are considered light ( $m_q \ll \mu_R$ ), and in which the remaining heavier quark flavors decouple from the theory. One may relate the coupling for the theory with  $n_f + 1$  light flavors to that with  $n_f$  flavors through an equation of the form

$$\alpha_s^{(n_f+1)}(\mu_R^2) = \alpha_s^{(n_f)}(\mu_R^2) \left( 1 + \sum_{n=1}^{\infty} \sum_{\ell=0}^n c_{n\ell} [\alpha_s^{(n_f)}(\mu_R^2)]^n \ln^\ell \frac{\mu_R^2}{m_h^2} \right), \quad (9.4)$$

where  $m_h$  is the mass of the  $(n_f+1)^{\text{th}}$  flavor, and the first few  $c_{n\ell}$  coefficients are  $c_{11} = \frac{1}{6\pi}$ ,  $c_{10} = 0$ ,  $c_{22} = c_{11}^2$ ,  $c_{21} = \frac{19}{24\pi^2}$ , and  $c_{20} = -\frac{11}{72\pi^2}$  when  $m_h$  is the  $\overline{\text{MS}}$  mass at scale  $m_h$  ( $c_{20} = \frac{7}{24\pi^2}$  when  $m_h$  is the pole mass — mass definitions are discussed below and in the review on “Quark Masses”). Terms up to  $c_{4\ell}$  are to be found in Refs. 16, 17. Numerically, when one chooses  $\mu_R = m_h$ , the matching is a modest effect, owing to the zero value for the  $c_{10}$  coefficient. Relations between  $n_f$  and  $(n_f+2)$  flavors where the two heavy flavors are close in mass are given to three loops in Ref. 18.

Working in an energy range where the number of flavors is taken constant, a simple exact analytic solution exists for Eq. (9.3) only if one neglects all but the  $b_0$  term, giving  $\alpha_s(\mu_R^2) = (b_0 \ln(\mu_R^2/\Lambda^2))^{-1}$ . Here  $\Lambda$  is a constant of integration, which corresponds to the scale where the perturbatively-defined coupling would diverge. Its value is indicative of the energy range where non-perturbative dynamics dominates. A convenient approximate analytic solution to the RGE that includes also the  $b_1$ ,  $b_2$ , and  $b_3$  terms is given by (see for example Ref. 19),

$$\alpha_s(\mu_R^2) \simeq \frac{1}{b_0 t} \left( 1 - \frac{b_1 \ln t}{b_0^2 t} + \frac{b_1^2 (\ln^2 t - \ln t - 1) + b_0 b_2}{b_0^4 t^2} - \frac{b_1^3 (\ln^3 t - \frac{5}{2} \ln^2 t - 2 \ln t + \frac{1}{2}) + 3b_0 b_1 b_2 \ln t - \frac{1}{2} b_0^2 b_3}{b_0^6 t^3} \right), \quad (9.5)$$

$t \equiv \ln \frac{\mu_R^2}{\Lambda^2},$

again parametrized in terms of a constant  $\Lambda$ . Note that Eq. (9.5) is one of several possible approximate 4-loop solutions for  $\alpha_s(\mu_R^2)$ , and that a value for  $\Lambda$  only defines  $\alpha_s(\mu_R^2)$  once one knows which particular

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approximation is being used. An alternative to the use of formulas such as Eq. (9.5) is to solve the RGE exactly, numerically (including the discontinuities, Eq. (9.4), at flavor thresholds). In such cases the quantity  $\Lambda$  is not defined at all. For these reasons, in determinations of the coupling, it has become standard practice to quote the value of  $\alpha_s$  at a given scale (typically the mass of the  $Z$  boson,  $M_Z$ ) rather than to quote a value for  $\Lambda$ .

The value of the coupling, as well as the exact forms of the  $b_2$ ,  $c_{10}$  (and higher-order) coefficients, depend on the renormalization scheme in which the coupling is defined, *i.e.* the convention used to subtract infinities in the context of renormalization. The coefficients given above hold for a coupling defined in the  $\overline{\text{MS}}$  scheme.

A discussion of determinations of the coupling and a graph illustrating its scale dependence (“running”) are to be found in Section 9.4. The RunDec package [20,21] is often used to calculate the evolution of the coupling. For a discussion of electroweak effects in the evolution of the QCD coupling, see Ref. 22 and references therein.

### 9.1.2. Quark masses :

Free quarks have never been observed, which is understood as a result of a long-distance, confining property of the strong QCD force: up, down, strange, charm, and bottom quarks all *hadronize*, *i.e.* become part of a meson or baryon, on a timescale  $\sim 1/\Lambda$ ; the top quark instead decays before it has time to hadronize. This means that the question of what one means by the quark mass is a complex one, which requires that one adopts a specific prescription. A perturbatively defined prescription is the pole mass,  $m_q$ , which corresponds to the position of the divergence of the propagator. This is close to one’s physical picture of mass. However, when relating it to observable quantities, it suffers from substantial non-perturbative ambiguities (see *e.g.* Ref. 23). An alternative is the  $\overline{\text{MS}}$  mass,  $\overline{m}_q(\mu_R^2)$ , which depends on the renormalization scale  $\mu_R$ .

Results for the masses of heavier quarks are often quoted either as the pole mass or as the  $\overline{\text{MS}}$  mass evaluated at a scale equal to the mass,  $\overline{m}_q(\overline{m}_q^2)$ ; light quark masses are often quoted in the  $\overline{\text{MS}}$  scheme at a scale  $\mu_R \sim 2 \text{ GeV}$ . The pole and  $\overline{\text{MS}}$  masses are related by a slowly converging series that starts  $m_q = \overline{m}_q(\overline{m}_q^2)(1 + \frac{4\alpha_s(\overline{m}_q^2)}{3\pi} + \mathcal{O}(\alpha_s^2))$ , while the scale-dependence of  $\overline{\text{MS}}$  masses is given by

$$\mu_R^2 \frac{d\overline{m}_q(\mu_R^2)}{d\mu_R^2} = \left[ -\frac{\alpha_s(\mu_R^2)}{\pi} + \mathcal{O}(\alpha_s^2) \right] \overline{m}_q(\mu_R^2). \quad (9.6)$$

More detailed discussion is to be found in a dedicated section of the *Review*, “Quark Masses.”

In perturbative QCD calculations of scattering processes, it is common to work in an approximation in which one neglects (*i.e.* sets to zero) the masses of all quarks whose mass is significantly smaller than the momentum transfer in the process.

## 9.2. Structure of QCD predictions

### 9.2.1. Fully inclusive cross sections :

The simplest observables in perturbative QCD are those that do not involve initial-state hadrons and that are fully inclusive with respect to details of the final state. One example is the total cross section for  $e^+e^- \rightarrow \text{hadrons}$  at center-of-mass energy  $Q$ , for which one can write

$$\frac{\sigma(e^+e^- \rightarrow \text{hadrons}, Q)}{\sigma(e^+e^- \rightarrow \mu^+\mu^-, Q)} \equiv R(Q) = R_{\text{EW}}(Q)(1 + \delta_{\text{QCD}}(Q)), \quad (9.7)$$

where  $R_{\text{EW}}(Q)$  is the purely electroweak prediction for the ratio and  $\delta_{\text{QCD}}(Q)$  is the correction due to QCD effects. To keep the discussion simple, we can restrict our attention to energies  $Q \ll M_Z$ , where the process is dominated by photon exchange ( $R_{\text{EW}} = 3 \sum_q e_q^2$ , neglecting finite-quark-mass corrections, where the  $e_q$  are the electric charges of the quarks),

$$\delta_{\text{QCD}}(Q) = \sum_{n=1}^{\infty} c_n \cdot \left( \frac{\alpha_s(Q^2)}{\pi} \right)^n + \mathcal{O}\left( \frac{\Lambda^4}{Q^4} \right). \quad (9.8)$$

The first four terms in the  $\alpha_s$  series expansion are then to be found in Ref. 24,

$$c_1 = 1, \quad c_2 = 1.9857 - 0.1152n_f, \quad (9.9a)$$

$$c_3 = -6.63694 - 1.20013n_f - 0.00518n_f^2 - 1.240\eta, \quad (9.9b)$$

$$c_4 = -156.61 + 18.775n_f - 0.7974n_f^2 + 0.0215n_f^3 + (17.828 - 0.575n_f)\eta, \quad (9.9c)$$

with  $\eta = (\sum e_q)^2 / (3 \sum e_q^2)$ . For corresponding expressions including also  $Z$  exchange and finite-quark-mass effects, see Refs. [25–27].

A related series holds also for the QCD corrections to the hadronic decay width of the  $\tau$  lepton, which essentially involves an integral of  $R(Q)$  over the allowed range of invariant masses of the hadronic part of the  $\tau$  decay (see *e.g.* Ref. 28). The series expansions for QCD corrections to Higgs-boson hadronic (partial) decay widths are summarized in Refs. 29, 30.

One characteristic feature of Eqs. (9.8) and (9.9) is that the coefficients of  $\alpha_s^n$  increase rapidly order by order: calculations in perturbative QCD tend to converge more slowly than would be expected based just on the size of  $\alpha_s^{\text{††}}$ . Another feature is the existence of an extra “power-correction” term  $\mathcal{O}(\Lambda^4/Q^4)$  in Eq. (9.8), which accounts for contributions that are fundamentally non-perturbative. All high-energy QCD predictions involve such corrections, though the exact power of  $\Lambda/Q$  depends on the observable. For many processes and observables, it is possible to introduce an operator product expansion and associate power suppressed terms with specific higher-dimension (non-perturbative) operators.

**Scale dependence.** In Eq. (9.8) the renormalization scale for  $\alpha_s$  has been chosen equal to  $Q$ . The result can also be expressed in terms of the coupling at an arbitrary renormalization scale  $\mu_R$ ,

$$\delta_{\text{QCD}}(Q) = \sum_{n=1}^{\infty} \bar{c}_n \left( \frac{\mu_R^2}{Q^2} \right) \cdot \left( \frac{\alpha_s(\mu_R^2)}{\pi} \right)^n + \mathcal{O}\left( \frac{\Lambda^4}{Q^4} \right), \quad (9.10)$$

where  $\bar{c}_1(\mu_R^2/Q^2) \equiv c_1$ ,  $\bar{c}_2(\mu_R^2/Q^2) = c_2 + \pi b_0 c_1 \ln(\mu_R^2/Q^2)$ ,  $\bar{c}_3(\mu_R^2/Q^2) = c_3 + (2b_0 c_2 \pi + b_1 c_1 \pi^2) \ln(\mu_R^2/Q^2) + b_0^2 c_1 \pi^2 \ln^2(\mu_R^2/Q^2)$ , *etc.* Given an infinite number of terms in the  $\alpha_s$  expansion, the  $\mu_R$  dependence of the  $\bar{c}_n(\mu_R^2/Q^2)$  coefficients will exactly cancel that of  $\alpha_s(\mu_R^2)$ , and the final result will be independent of the choice of  $\mu_R$ : physical observables do not depend on unphysical scales.\*\*

With just terms up to some finite  $n = N$ , a residual  $\mu_R$  dependence will remain, which implies an uncertainty on the prediction of  $R(Q)$  due to the arbitrariness of the scale choice. This uncertainty will be  $\mathcal{O}(\alpha_s^{N+1})$ , *i.e.* of the same order as the neglected terms. For this reason it is customary to use QCD predictions’ scale dependence as an estimate of the uncertainties due to neglected terms. One usually takes a central value for  $\mu_R \sim Q$ , in order to avoid the poor convergence of the perturbative series that results from the large  $\ln^{n-1}(\mu_R^2/Q^2)$  terms in the  $\bar{c}_n$  coefficients when  $\mu_R \ll Q$  or  $\mu_R \gg Q$ . Uncertainties are then commonly determined by varying  $\mu_R$  by a factor of two up and down around the central scale choice, as discussed in more detail below in Section 9.2.4.

<sup>††</sup> The situation is significantly worse near thresholds, *e.g.* the  $t\bar{t}$  production threshold. An overview of some of the methods used in such cases is to be found for example in Ref. 31.

\*\* There is an important caveat to this statement: at sufficiently high orders, perturbative series generally suffer from “renormalon” divergences  $\alpha_s^n n!$  (reviewed in Ref. 23). This phenomenon is not usually visible with the limited number of perturbative terms available today. However it is closely connected with non-perturbative contributions and sets a limit on the possible precision of perturbative predictions. The cancellation of scale dependence will also ultimately be affected by this renormalon-induced breakdown of perturbation theory.

### 9.2.2. Processes with initial-state hadrons :

**Deep Inelastic Scattering.** To illustrate the key features of QCD cross sections in processes with initial-state hadrons, let us consider deep-inelastic scattering (DIS),  $ep \rightarrow e + X$ , where an electron  $e$  with four-momentum  $k$  emits a highly off-shell photon (momentum  $q$ ) that interacts with the proton (momentum  $p$ ). For photon virtualities  $Q^2 \equiv -q^2$  far above the squared proton mass (but far below the  $Z$  mass), the differential cross section in terms of the kinematic variables  $Q^2$ ,  $x = Q^2/(2p \cdot q)$  and  $y = (q \cdot p)/(k \cdot p)$  is

$$\frac{d^2\sigma}{dx dQ^2} = \frac{4\pi\alpha}{2xQ^4} \left[ (1 + (1-y)^2)F_2(x, Q^2) - y^2 F_L(x, Q^2) \right], \quad (9.11)$$

where  $\alpha$  is the electromagnetic coupling and  $F_2(x, Q^2)$  and  $F_L(x, Q^2)$  are proton structure functions, which encode the interaction between the photon (in given polarization states) and the proton. In the presence of parity-violating interactions (*e.g.*  $\nu p$  scattering) an additional  $F_3$  structure function is present. For an extended review, including equations for the full electroweak and polarized cases, see Sec. 19 of this *Review*.

Structure functions are not calculable in perturbative QCD, nor is any other cross section that involves initial-state hadrons. To zeroth order in  $\alpha_s$ , the structure functions are given directly in terms of non-perturbative parton (quark or gluon) distribution functions (PDFs),

$$F_2(x, Q^2) = x \sum_q e_q^2 f_{q/p}(x), \quad F_L(x, Q^2) = 0, \quad (9.12)$$

where  $f_{q/p}(x)$  is the PDF for quarks of type  $q$  inside the proton, *i.e.* the number density of quarks of type  $q$  inside a fast-moving proton that carry a fraction  $x$  of its longitudinal momentum (the quark flavor index  $q$ , here, is not to be confused with the photon momentum  $q$  in the lines preceding Eq. (9.11)). PDFs are non-perturbative, and only just starting to be extracted in lattice QCD in a phenomenologically relevant way [32]. Accordingly, for all practical uses, they are determined from data (*cf.* Sec. 19 of this *Review* and also Ref. 33).

The above result, with PDFs  $f_{q/p}(x)$  that are independent of the scale  $Q$ , corresponds to the “quark-parton model” picture in which the photon interacts with point-like free quarks, or equivalently, one has incoherent elastic scattering between the electron and individual constituents of the proton. As a consequence, in this picture also  $F_2$  and  $F_L$  are independent of  $Q$  [34]. When including higher orders in pQCD, Eq. (9.12) becomes

$$F_2(x, Q^2) = x \sum_{n=0}^{\infty} \frac{\alpha_s^n(\mu_R^2)}{(2\pi)^n} \sum_{i=q,g} \int_x^1 \frac{dz}{z} C_{2,i}^{(n)}(z, Q^2, \mu_R^2, \mu_F^2) f_{i/p}\left(\frac{x}{z}, \mu_F^2\right) + \mathcal{O}\left(\frac{\Lambda^2}{Q^2}\right). \quad (9.13)$$

Just as in Eq. (9.10), we have a series in powers of  $\alpha_s(\mu_R^2)$ , each term involving a coefficient  $C_{2,i}^{(n)}$  that can be calculated using Feynman graphs. An important difference is the additional integral over  $z$ . The parton that comes from the proton can emit a gluon before it interacts with the photon. As a result, the  $C_{2,i}^{(n)}$  coefficients are functions that depend on the ratio,  $z$ , of the parton’s momentum before and after the gluon emission, and one must integrate over that ratio. For the electromagnetic component of DIS with light quarks and gluons, the zeroth order coefficient functions are  $C_{2,q}^{(0)} = e_q^2 \delta(1-z)$  and  $C_{2,g}^{(0)} = 0$ , and corrections are known up to  $\mathcal{O}(\alpha_s^3)$  (next-to-next-next-to-leading order, N<sup>3</sup>LO) [35]. For weak currents they are known fully to  $\alpha_s^2$  (next-to-next-to-leading order, NNLO) [36], with substantial results known also at N<sup>3</sup>LO [37]. For heavy quark production they are known to  $\mathcal{O}(\alpha_s^2)$  [38] (next-to-leading order, NLO, insofar as the series starts at  $\mathcal{O}(\alpha_s)$ ), with ongoing work towards NNLO summarized in Ref. 39.

The majority of the emissions that modify a parton’s momentum are collinear (parallel) to that parton, and don’t depend on the fact that the parton is destined to interact with a photon. It is natural to view these emissions as modifying the proton’s structure rather

than being part of the coefficient function for the parton’s interaction with the photon. Technically, one uses a procedure known as *collinear factorization* to give a well-defined meaning to this distinction, most commonly through the  $\overline{\text{MS}}$  factorization scheme, defined in the context of dimensional regularization. The  $\overline{\text{MS}}$  factorization scheme involves an arbitrary choice of *factorization scale*,  $\mu_F$ , whose meaning can be understood roughly as follows: emissions with transverse momenta above  $\mu_F$  are included in the  $C_{2,q}^{(n)}(z, Q^2, \mu_R^2, \mu_F^2)$ ; emissions with transverse momenta below  $\mu_F$  are accounted for within the PDFs,  $f_{i/p}(x, \mu_F^2)$ . While collinear factorization is generally believed to be valid for suitable (sufficiently inclusive) observables in processes with hard scales, Ref. 40, which reviews the factorization proofs in detail, is cautious in the statements it makes about their exhaustivity, notably for the hadron-collider processes that we shall discuss below. Further discussion is to be found in Refs. 41,42.

The PDFs’ resulting dependence on  $\mu_F$  is described by the Dokshitzer-Gribov-Lipatov-Altarelli-Parisi (DGLAP) equations [43], which to leading order (LO) read\*

$$\mu_F^2 \frac{\partial f_{i/p}(x, \mu_F^2)}{\partial \mu_F^2} = \sum_j \frac{\alpha_s(\mu_F^2)}{2\pi} \int_x^1 \frac{dz}{z} P_{i \leftarrow j}^{(1)}(z) f_{j/p}\left(\frac{x}{z}, \mu_F^2\right), \quad (9.14)$$

with, for example,  $P_{q \leftarrow g}^{(1)}(z) = T_R(z^2 + (1-z)^2)$ . The other LO splitting functions are listed in Sec. 19 of this *Review*, while results up to NLO,  $\alpha_s^2$ , and NNLO,  $\alpha_s^3$ , are given in Refs. 44 and 45 respectively. Beyond LO, the coefficient functions are also  $\mu_F$  dependent, for example  $C_{2,i}^{(1)}(x, Q^2, \mu_R^2, \mu_F^2) = C_{2,i}^{(1)}(x, Q^2, \mu_R^2, Q^2) - \ln\left(\frac{\mu_F^2}{Q^2}\right) \sum_j \int_x^1 \frac{dz}{z} C_{2,j}^{(0)}\left(\frac{x}{z}\right) P_{j \leftarrow i}^{(1)}(z)$ .

As with the renormalization scale, the choice of factorization scale is arbitrary, but if one has an infinite number of terms in the perturbative series, the  $\mu_F$ -dependences of the coefficient functions and PDFs will compensate each other fully. Given only  $N$  terms of the series, a residual  $\mathcal{O}(\alpha_s^{N+1})$  uncertainty is associated with the ambiguity in the choice of  $\mu_F$ . As with  $\mu_R$ , varying  $\mu_F$  provides an input in estimating uncertainties on predictions. In inclusive DIS predictions, the default choice for the scales is usually  $\mu_R = \mu_F = Q$ .

As is the case for the running coupling, in DGLAP evolution one can introduce flavor thresholds near the heavy quark masses: below a given heavy quark’s mass, that quark is not considered to be part of the proton’s structure, while above it is considered to be part of the proton’s structure and evolves with massless DGLAP splitting kernels. With appropriate parton distribution matching terms at threshold, such a variable flavor number scheme (VFNS), when used with massless coefficient functions, gives the full heavy-quark contributions at high  $Q^2$  scales. For scales near the threshold, it is instead necessary to appropriately adapt the standard massive coefficient functions to account for the heavy-quark contribution already included in the PDFs [46,47,48].

**Hadron-hadron collisions.** The extension to processes with two initial-state hadrons can be illustrated with the example of the total (inclusive) cross section for  $W$  boson production in collisions of hadrons  $h_1$  and  $h_2$ , which can be written as

$$\begin{aligned} & \sigma(h_1 h_2 \rightarrow W + X) \\ &= \sum_{n=0}^{\infty} \alpha_s^n(\mu_R^2) \sum_{i,j} \int dx_1 dx_2 f_{i/h_1}(x_1, \mu_F^2) f_{j/h_2}(x_2, \mu_F^2) \\ & \times \hat{\sigma}_{ij \rightarrow W+X}^{(n)}(x_1 x_2 s, \mu_R^2, \mu_F^2) + \mathcal{O}\left(\frac{\Lambda^2}{M_W^4}\right), \end{aligned} \quad (9.15)$$

\* LO is generally taken to mean the lowest order at which a quantity is non-zero. This definition is nearly always unambiguous, the one major exception being for the case of the hadronic branching ratio of virtual photons,  $Z$ ,  $\tau$ , *etc.*, for which two conventions exist: LO can either mean the lowest order that contributes to the hadronic branching fraction, *i.e.* the term “1” in Eq. (9.7); or it can mean the lowest order at which the hadronic branching ratio becomes sensitive to the coupling,  $n = 1$  in Eq. (9.8), as is relevant when extracting the value of the coupling from a measurement of the branching ratio. Because of this ambiguity, we avoid use of the term “LO” in that context.

where  $s$  is the squared center-of-mass energy of the collision. At LO,  $n = 0$ , the hard (partonic) cross section  $\hat{\sigma}_{ij \rightarrow W+X}^{(0)}(x_1 x_2 s, \mu_R^2, \mu_F^2)$  is simply proportional to  $\delta(x_1 x_2 s - M_W^2)$ , in the narrow  $W$ -boson width approximation (see Sec. 49 of this *Review* for detailed expressions for this and other hard scattering cross sections). It is non-zero only for choices of  $i, j$  that can directly give a  $W$ , such as  $i = u, j = \bar{d}$ . At higher orders,  $n \geq 1$ , new partonic channels contribute, such as  $gq$ , and there is no restriction  $x_1 x_2 s = M_W^2$ .

Equation (9.15) involves a collinear factorization between the hard cross section and the PDFs, just like Eq. (9.13). As long as the same factorization scheme is used in DIS and  $pp$  or  $p\bar{p}$  (usually the  $\overline{\text{MS}}$  scheme), then PDFs extracted in DIS can be directly used in  $pp$  and  $p\bar{p}$  predictions [49,40] (with the anti-quark distributions in an anti-proton being the same as the quark distributions in a proton).

Fully inclusive hard cross sections are known to NNLO, *i.e.* corrections up to relative order  $\alpha_s^2$ , for Drell-Yan (DY) lepton-pair and vector-boson production [50,51], Higgs-boson production in association with a vector boson [52], Higgs-boson production via vector-boson fusion [53] (in an approximation that factorizes the production of the two vector bosons), Higgs-pair production [54], top-antitop production [55] and vector-boson pair production [56,57].<sup>†</sup> Recently, inclusive Higgs production through gluon fusion was calculated at N<sup>3</sup>LO [58]. A discussion of many other inclusive Higgs results is to be found in Ref. 59.

**Photoproduction.**  $\gamma p$  (and  $\gamma\gamma$ ) collisions are similar to  $pp$  collisions, with the subtlety that the photon can behave in two ways: there is “direct” photoproduction, in which the photon behaves as a point-like particle and takes part directly in the hard collision, with hard subprocesses such as  $\gamma g \rightarrow q\bar{q}$ ; there is also resolved photoproduction, in which the photon behaves like a hadron, with non-perturbative partonic substructure and a corresponding PDF for its quark and gluon content,  $f_{i/\gamma}(x, Q^2)$ .

While useful to understand the general structure of  $\gamma p$  collisions, the distinction between direct and resolved photoproduction is not well defined beyond leading order, as discussed for example in Ref. 60.

**The high-energy (BFKL) limit.** In situations in which the total center-of-mass energy  $\sqrt{s}$  is much larger than all other momentum-transfer scales in the problem (*e.g.*  $Q$  in DIS,  $m_b$  for  $b\bar{b}$  production in  $pp$  collisions, *etc.*), each power of  $\alpha_s$  beyond LO can be accompanied by a power of  $\ln(s/Q^2)$  (or  $\ln(s/m_b^2)$ , *etc.*). This is variously referred to as the high-energy, small- $x$  or Balitsky-Fadin-Kuraev-Lipatov (BFKL) limit [61–63]. Currently it is possible to account for the dominant and first subdominant [64,65] power of  $\ln s$  at each order of  $\alpha_s$ , and also to estimate further subdominant contributions that are numerically large (see Refs. 66–69 and references therein). Progress towards NNLO is discussed in Refs. 70,71.

Physically, the summation of all orders in  $\alpha_s$  can be understood as leading to a growth with  $s$  of the gluon density in the proton. At sufficiently high energies this implies non-linear effects (commonly referred to as parton saturation), whose treatment has been the subject of intense study (see for example Refs. 72, 73 and references thereto). Note that it is not straightforward to relate these results to the genuinely non-perturbative total, elastic and diffractive cross sections for hadron-hadron scattering (experimental results for which are summarized in section 51 of this *Review*).

### 9.2.3. Non fully inclusive cross sections :

QCD final states always consist of hadrons, while perturbative QCD calculations deal with partons. Physically, an energetic parton fragments (“showers”) into many further partons, which then, on later timescales, undergo a transition to hadrons (“hadronization”). Fixed-order perturbation theory captures only a small part of these dynamics.

This does not matter for the fully inclusive cross sections discussed above: the showering and hadronization stages are approximately unitary, *i.e.* they do not substantially change the overall probability

<sup>†</sup> Processes with jets or photons in the final state have divergent cross sections unless one places cut on the jet or photon momentum. Accordingly they are discussed below in Section 9.2.3.2.

of hard scattering, because they occur long after it has taken place (they introduce at most a correction proportional to a power of the ratio of timescales involved, *i.e.* a power of  $\Lambda/Q$ , where  $Q$  is the hard scattering scale).

Less inclusive measurements, in contrast, may be affected by the extra dynamics. For those sensitive just to the main directions of energy flow (jet rates, event shapes, *cf.* Sec. 9.3.1) fixed order perturbation theory is often still adequate, because showering and hadronization don’t substantially change the overall energy flow. This means that one can make a prediction using just a small number of partons, which should correspond well to a measurement of the same observable carried out on hadrons. For observables that instead depend on distributions of individual hadrons (which, *e.g.*, are the inputs to detector simulations), it is mandatory to account for showering and hadronization. The range of predictive techniques available for QCD final states reflects this diversity of needs of different measurements.

While illustrating the different methods, we shall for simplicity mainly use expressions that hold for  $e^+e^-$  scattering. The extension to cases with initial-state partons will be mostly straightforward (space constraints unfortunately prevent us from addressing diffraction and exclusive hadron-production processes; extensive discussion is to be found in Refs. 74, 75).

#### 9.2.3.1. Soft and collinear limits:

Before examining specific predictive methods, it is useful to be aware of a general property of QCD matrix elements in the soft and collinear limits. Consider a squared tree-level matrix element  $|M_n^2(p_1, \dots, p_n)|$  for the process  $e^+e^- \rightarrow n$  partons with momenta  $p_1, \dots, p_n$ , and a corresponding phase-space integration measure  $d\Phi_n$ . If particle  $n$  is a gluon, and additionally it becomes collinear (parallel) to another particle  $i$  and its momentum tends to zero (it becomes “soft”), the matrix element simplifies as follows,

$$\lim_{\theta_{in} \rightarrow 0, E_n \rightarrow 0} d\Phi_n |M_n^2(p_1, \dots, p_n)| = d\Phi_{n-1} |M_{n-1}^2(p_1, \dots, p_{n-1})| \frac{\alpha_s C_i}{\pi} \frac{d\theta_{in}^2}{\theta_{in}^2} \frac{dE_n}{E_n}, \quad (9.16)$$

where  $C_i = C_F$  ( $C_A$ ) if  $i$  is a quark (gluon). This formula has non-integrable divergences both for the inter-parton angle  $\theta_{in} \rightarrow 0$  and for the gluon energy  $E_n \rightarrow 0$ , which are mirrored also in the structure of divergences in loop diagrams. These divergences are important for at least two reasons: firstly, they govern the typical structure of events (inducing many emissions either with low energy or at small angle with respect to hard partons); secondly, they will determine which observables can be calculated within perturbative QCD.

#### 9.2.3.2. Fixed-order predictions:

Let us consider an observable  $\mathcal{O}$  that is a function  $\mathcal{O}_n(p_1, \dots, p_n)$  of the four-momenta of the  $n$  final-state particles in an event (whether partons or hadrons). In what follows, we shall consider the cross section for events weighted with the value of the observable,  $\sigma_{\mathcal{O}}$ . As examples, if  $\mathcal{O}_n \equiv 1$  for all  $n$ , then  $\sigma_{\mathcal{O}}$  is just the total cross section; if  $\mathcal{O}_n \equiv \hat{\tau}(p_1, \dots, p_n)$  where  $\hat{\tau}$  is the value of the Thrust for that event (see Sec. 9.3.1.2), then the average value of the Thrust is  $\langle \tau \rangle = \sigma_{\mathcal{O}}/\sigma_{\text{tot}}$ ; if  $\mathcal{O}_n \equiv \delta(\tau - \hat{\tau}(p_1, \dots, p_n))$  then one gets the differential cross section as a function of the Thrust,  $\sigma_{\mathcal{O}} \equiv d\sigma/d\tau$ .

In the expressions below, we shall omit to write the non-perturbative power correction term, which for most common observables is proportional to a single power of  $\Lambda/Q$ .

**LO.** If the observable  $\mathcal{O}$  is non-zero only for events with at least  $n$  final-state particles, then the LO QCD prediction for the weighted cross section in  $e^+e^-$  annihilation is

$$\sigma_{\mathcal{O},LO} = \alpha_s^{n-2}(\mu_R^2) \int d\Phi_n |M_n^2(p_1, \dots, p_n)| \mathcal{O}_n(p_1, \dots, p_n), \quad (9.17)$$

where the squared tree-level matrix element,  $|M_n^2(p_1, \dots, p_n)|$ , includes relevant symmetry factors, has been summed over all subprocesses (*e.g.*  $e^+e^- \rightarrow q\bar{q}q\bar{q}$ ,  $e^+e^- \rightarrow q\bar{q}gg$ ) and has had all factors of  $\alpha_s$

extracted in front. In processes other than  $e^+e^-$  collisions, the center-of-mass energy of the LO process is generally not fixed, and so the powers of the coupling are often brought inside the integrals, with the scale  $\mu_R$  chosen event by event, as a function of the event kinematics.

Other than in the simplest cases (see the review on Cross Sections in this *Review*), the matrix elements in Eq. (9.17) are usually calculated automatically with programs such as CompHEP [76], MadGraph [77], Alpgen [78], Comix/Sherpa [79], and Helac/Phegas [80]. Some of these (CompHEP, MadGraph) use formulas obtained from direct evaluations of Feynman diagrams. Others (Alpgen, Helac/Phegas and Comix/Sherpa) use methods designed to be particularly efficient at high multiplicities, such as Berends-Giele recursion [81], which builds up amplitudes for complex processes from simpler ones (see also the reviews and discussion in Refs. [82–84]).

The phase-space integration is usually carried out by Monte Carlo sampling, in order to deal with the sometimes complicated cuts that are used in corresponding experimental measurements. Because of the divergences in the matrix element, Eq. (9.16), the integral converges only if the observable vanishes for kinematic configurations in which one of the  $n$  particles is arbitrarily soft or it is collinear to another particle. As an example, the cross section for producing any configuration of  $n$  partons will lead to an infinite integral, whereas a finite result will be obtained for the cross section for producing  $n$  deposits of energy (or jets, see Sec. 9.3.1.1), each above some energy threshold and well separated from each other in angle.

LO calculations can be carried out for  $2 \rightarrow n$  processes with  $n \lesssim 6-10$ . The exact upper limit depends on the process, the method used to evaluate the matrix elements (recursive methods are more efficient), and the extent to which the phase-space integration can be optimized to work around the large variations in the values of the matrix elements.

**NLO.** Given an observable that is non-zero starting from  $n$  final-state particles, its prediction at NLO involves supplementing the LO result, Eq. (9.17), with the  $2 \rightarrow (n+1)$ -particle squared tree-level matrix element ( $|M_{n+1}^2|$ ), and the interference of an  $2 \rightarrow n$  tree-level and  $2 \rightarrow n$  1-loop amplitude ( $2\text{Re}(M_n M_{n,1\text{-loop}}^*)$ ),

$$\begin{aligned} \sigma_{\mathcal{O}}^{\text{NLO}} = & \sigma_{\mathcal{O}}^{\text{LO}} + \alpha_s^{n-1}(\mu_R^2) \int d\Phi_{n+1} |M_{n+1}^2(p_1, \dots, p_{n+1})| \mathcal{O}_{n+1}(p_1, \dots, p_{n+1}) \\ & + \alpha_s^{n-1}(\mu_R^2) \int d\Phi_n 2\text{Re} [ M_n(p_1, \dots, p_n) M_{n,1\text{-loop}}^*(p_1, \dots, p_n) ] \\ & \times \mathcal{O}_n(p_1, \dots, p_n) . \end{aligned} \quad (9.18)$$

Relative to LO calculations, two important issues appear in the NLO calculations. Firstly, the extra complexity of loop-calculations relative to tree-level calculations means that their automation has been achieved only in recent years (see below). Secondly, loop amplitudes are infinite in 4 dimensions, while tree-level amplitudes are finite, but their *integrals* are infinite, due to the divergences of Eq. (9.16). These two sources of infinities have the same soft and collinear origins and cancel after the integration only if the observable  $\mathcal{O}$  satisfies the property of infrared and collinear safety,

$$\begin{aligned} \mathcal{O}_{n+1}(p_1, \dots, p_s, \dots, p_n) & \rightarrow \mathcal{O}_n(p_1, \dots, p_n) & \text{if } p_s \rightarrow 0 \\ \mathcal{O}_{n+1}(p_1, \dots, p_a, p_b, \dots, p_n) & \rightarrow \mathcal{O}_n(p_1, \dots, p_a + p_b, \dots, p_n) \\ & \text{if } p_a \parallel p_b . \end{aligned} \quad (9.19)$$

Examples of infrared-safe quantities include event-shape distributions and jet cross sections (with appropriate jet algorithms, see below). Unsafe quantities include the distribution of the momentum of the hardest QCD particle (which is not conserved under collinear splitting), observables that require the complete absence of radiation in some region of phase space (*e.g.* rapidity gaps or 100% isolation cuts, which are affected by soft emissions), or the particle multiplicity (affected by both soft and collinear emissions). The non-cancellation of divergences at NLO due to infrared or collinear unsafety compromises the usefulness not only of the NLO calculation, but also that of a

LO calculation, since LO is only an acceptable approximation if one can prove that higher-order terms are smaller. Infrared and collinear unsafety usually also imply large non-perturbative effects.

As with LO calculations, the phase-space integrals in Eq. (9.18) are usually carried out by Monte Carlo integration, so as to facilitate the study of arbitrary observables. Various methods exist to obtain numerically efficient cancellation among the different infinities. These include notably dipole [85], FKS [86] and antenna [87] subtraction.

NLO calculations exist for a wide range of processes. Historically, many calculations have been performed process by process and are available in dedicated packages, among them NLOJet++ [88] for  $e^+e^-$ , DIS, and hadron-hadron processes involving just light partons in the final state, MCFM [89] for hadron-hadron processes with Higgs or vector bosons and/or heavy quarks in the final state, VBFNLO for vector-boson fusion, di- and tri-boson processes [90], and the Phox family [91] for processes with photons in the final state. Many of these programs are still widely used today.

Recent years have seen very active development of automated NLO calculational tools, and a number of programs are available publicly: Madgraph5\_aMC@NLO [77] and Helac-NLO [92] provide full frameworks for NLO calculations; GoSam [93], Njet [94], OpenLoops [95] calculate just the 1-loop part and are typically interfaced with an external tool such as Sherpa [96] for combination with the appropriate tree-level amplitudes. Other tools such as BlackHat [97] and Recola [98] are not currently available publicly, though in the case of the former many of its results can be accessed in the form of ntuples [99] to which a range of cuts, and histogramming options, as well as PDF and scale-changes, can be applied *a posteriori*; an alternative approach for *a posteriori* PDF and scale change represents NLO (or NNLO) results, for a given set of cuts and binning, as an effective coefficient function on a grid in parton momentum fractions and factorization scales [100–103].

In some cases the above programs (or development versions of them) can be used to calculate also NLO electroweak or beyond-standard-model corrections [104–107]. Electroweak corrections are especially important for transverse momenta significantly above the  $W$  and  $Z$  masses, because they are enhanced by two powers of  $\ln p_t/M_W$  for each power of the electroweak coupling.

The above tools rely in part on a wide array of developments reviewed in Refs. 83,108. Examples of the most complex processes for which NLO QCD corrections have been obtained so far include  $e^+e^- \rightarrow 7$  jets [109],  $pp \rightarrow W + 5$  jets [110] and  $pp \rightarrow 5$  jets [111].

**NNLO.** Conceptually, NNLO and NLO calculations are similar, except that one must add a further order in  $\alpha_s$ , consisting of: the squared  $(n+2)$ -parton tree-level amplitude, the interference of the  $(n+1)$ -parton tree-level and 1-loop amplitudes, the interference of the  $n$ -parton tree-level and 2-loop amplitudes, and the squared  $n$ -parton 1-loop amplitude.

Each of these elements involves large numbers of soft and collinear divergences, satisfying relations analogous to Eq. (9.16) that now involve multiple collinear or soft particles and higher loop orders (see *e.g.* Refs. [112–114]). Arranging for the cancellation of the divergences after numerical Monte Carlo integration has been one of the significant challenges of NNLO calculations, as has the determination of the relevant 2-loop amplitudes. For the cancellations of divergences a wide range of methods has been developed. Some of them [115–119] retain the approach, inherent in NLO methods, of directly combining the separate loop and tree-level amplitudes. Others combine a suitably chosen, partially inclusive  $2 \rightarrow n$  NNLO calculation with a fully differential  $2 \rightarrow n+1$  NLO calculation [120–123].

Quite a number of processes have been calculated differentially at NNLO so far. The state of the art for  $e^+e^-$  collisions is  $e^+e^- \rightarrow 3$  jets [124–126]. For hadron colliders, all  $2 \rightarrow 1$  processes are known, specifically vector boson [127,128] and Higgs boson production [129,120]. For most of the above calculations there exist public codes (EERAD3 for  $e^+e^-$ , DYNLO and FEWZ for  $W$  and  $Z$  production, fehipro and HNNLO for Higgs production), links to which are to be found among the above references. Substantial progress has been made in the past couple of years for hadron-collider  $2 \rightarrow 2$  processes, with calculations having been



performed for nearly all relevant processes:  $HH$  [54],  $WH$  [130] and  $ZH$  [131],  $ZZ$  [57] and  $WW$  [56],  $\gamma\gamma$  [132],  $Z\gamma$  [133] and  $W\gamma$  [134],  $H$ +jet [135,136],  $W$ +jet [121] and  $Z$ +jet [137] (the latter in a leading-color approximation for the dominant partonic channels),  $t$ -channel single-top [138] (in a so-called “structure-function” approximation) and  $t\bar{t}$  production [139]. Of notable interest, but still with only partial NNLO results, is dijet production [140]. One  $2 \rightarrow 3$  process is known at NNLO, Higgs production through vector-boson fusion, using an approximation in which the two underlying DIS-like  $q \rightarrow qV$  scatterings are factorised [123].

### 9.2.3.3. Resummation:

Many experimental measurements place tight constraints on emissions in the final state. For example, in  $e^+e^-$  events, that (one minus) the Thrust should be less than some value  $\tau \ll 1$ , or in  $pp \rightarrow Z$  events that the  $Z$ -boson transverse momentum should be much smaller than its mass,  $p_t^Z \ll M_Z$ . A further example is the production of heavy particles or jets near threshold (so that little energy is left over for real emissions) in DIS and  $pp$  collisions.

In such cases, the constraint vetoes a significant part of the integral over the soft and collinear divergence of Eq. (9.16). As a result, there is only a partial cancellation between real emission terms (subject to the constraint) and loop (virtual) contributions (not subject to the constraint), causing each order of  $\alpha_s$  to be accompanied by a large coefficient  $\sim L^2$ , where *e.g.*  $L = \ln \tau$  or  $L = \ln(M_Z/p_t^Z)$ . One ends up with a perturbative series whose terms go as  $\sim (\alpha_s L^2)^n$ . It is not uncommon that  $\alpha_s L^2 \gg 1$ , so that the perturbative series converges very poorly if at all.\*\* In such cases one may carry out a “resummation,” which accounts for the dominant logarithmically enhanced terms to all orders in  $\alpha_s$ , by making use of known properties of matrix elements for multiple soft and collinear emissions, and of the all-orders properties of the divergent parts of virtual corrections, following original works such as Refs. 141–150 and also through soft-collinear effective theory [151,152] (*cf.* also the section on “Heavy-Quark and Soft-Collinear Effective Theory” in this *Review*, as well as Ref. 153).

For cases with double logarithmic enhancements (two powers of logarithm per power of  $\alpha_s$ ), there are two classification schemes for resummation accuracy. Writing the cross section including the constraint as  $\sigma(L)$  and the unconstrained (total) cross section as  $\sigma_{\text{tot}}$ , the series expansion takes the form

$$\sigma(L) \simeq \sigma_{\text{tot}} \sum_{n=0}^{\infty} \sum_{k=0}^{2n} R_{nk} \alpha_s^n (\mu_R^2)^k L^k, \quad L \gg 1 \quad (9.20)$$

and leading log (LL) resummation means that one accounts for all terms with  $k = 2n$ , next-to-leading-log (NLL) includes additionally all terms with  $k = 2n - 1$ , *etc.* Often  $\sigma(L)$  (or its Fourier or Mellin transform) *exponentiates*<sup>†</sup>,

$$\sigma(L) \simeq \sigma_{\text{tot}} \exp \left[ \sum_{n=1}^{\infty} \sum_{k=0}^{n+1} G_{nk} \alpha_s^n (\mu_R^2)^k L^k \right], \quad L \gg 1, \quad (9.21)$$

where one notes the different upper limit on  $k$  ( $\leq n+1$ ) compared to Eq. (9.20). This is a more powerful form of resummation: the  $G_{12}$  term alone reproduces the full LL series in Eq. (9.20). With the form Eq. (9.21) one still uses the nomenclature LL, but this now means that all terms with  $k = n+1$  are included, and NLL implies all terms with  $k = n$ , *etc.*

\*\* To be precise one should be aware of two causes of the divergence of perturbative series. That which interests us here is associated with the presence of a new large parameter (*e.g.* ratio of scales). It is distinct from the “renormalon” induced factorial divergences of perturbation theory that were discussed above.

† Whether or not this happens depends on the quantity being resummed. A classic example involves jet rates in  $e^+e^-$  collisions as a function of a jet-resolution parameter  $y_{\text{cut}}$ . The logarithms of  $1/y_{\text{cut}}$  exponentiate for the  $k_t$  (Durham) jet algorithm [154], but not [155] for the JADE algorithm [156] (both are discussed below in Sec. 9.3.1.1).

For a large number of observables, NLL resummations are available in the sense of Eq. (9.21) (see Refs. 157–159 and references therein). NNLL has been achieved for the DY and Higgs-boson  $p_t$  distributions [160–163] (also available in the CuTe [164], HRes [165] and ResBos [166] families of programs and also differentially in vector-boson decay products [167]) and related variables [168], for the  $p_t$  of vector-boson pairs [169], for the back-to-back energy-energy correlation in  $e^+e^-$  [170], the jet broadening in  $e^+e^-$  collisions [171], the jet-veto survival probability in Higgs and  $Z$  boson production in  $pp$  collisions [172], an event-shape type observable known as the beam Thrust [173], hadron-collider jet masses in specific limits [174] (see also Ref. 175), the production of top anti-top pairs near threshold [176–178] (and references therein), and high- $p_t$   $W$  and  $Z$  production [179]. Automation of NNLL jet-veto resummations for different processes has been achieved in Ref. 180 (*cf.* also the NLL automation in Ref. 181), while automation for a certain class of  $e^+e^-$  observables has been achieved in Ref. 182. The parts believed to be dominant in the N<sup>3</sup>LL resummation are available for the Thrust variable,  $C$ -parameter and heavy-jet mass in  $e^+e^-$  annihilations [183–185] (confirmed for Thrust at NNLL in Ref. 186), and full N<sup>3</sup>LL has been achieved for Higgs- and vector-boson production near threshold [187]. An extensive discussion of jet masses for heavy-quark induced jets has been given in Ref. 188. Recently, there has also been progress in resummed calculations for jet substructure, whose observables involve more complicated definitions than is the case for standard resummations [189–193]. The inputs and methods involved in these various calculations are somewhat too diverse to discuss in detail here, so we recommend that the interested reader consult the original references for further details.

### 9.2.3.4. Fragmentation functions:

Since the parton-hadron transition is non-perturbative, it is not possible to perturbatively calculate quantities such as the energy-spectra of specific hadrons in high-energy collisions. However, one can factorize perturbative and non-perturbative contributions via the concept of fragmentation functions. These are the final-state analogue of the parton distribution functions that are used for initial-state hadrons. Like parton distribution functions, they depend on a (fragmentation) factorization scale and satisfy a DGLAP evolution equation.

It should be added that if one ignores the non-perturbative difficulties and just calculates the energy and angular spectrum of partons in perturbative QCD with some low cutoff scale  $\sim \Lambda$  (using resummation to sum large logarithms of  $\sqrt{s}/\Lambda$ ), then this reproduces many features of the corresponding hadron spectra [194]. This is often taken to suggest that hadronization is “local”, in the sense it mainly involves partons that are close both in position and in momentum.

Section 20 of this *Review* provides further information (and references) on these topics, including also the question of heavy-quark fragmentation.

### 9.2.3.5. Parton-shower Monte Carlo generators:

Parton-shower Monte Carlo (MC) event generators like PYTHIA [195–197], HERWIG [198–200] and SHERPA [96] provide fully exclusive simulations of QCD events.<sup>†</sup> Because they provide access to “hadron-level” events, they are a crucial tool for all applications that involve simulating the response of detectors to QCD events. Here we give only a brief outline of how they work and refer the reader to Sec. 41 and Ref. 202 for a full overview.

The MC generation of an event involves several stages. It starts with the random generation of the kinematics and partonic channels of whatever *hard scattering process* the user has requested at some high scale  $Q_0$  (for complex processes, this may be carried out by an external program). This is followed by a *parton shower*, usually based on the successive random generation of gluon emissions (or  $g \rightarrow q\bar{q}$  splittings). Each is generated at a scale lower than the previous emission, following a (soft and collinear resummed) perturbative QCD

† The program ARIADNE [201] has also been widely used for simulating  $e^+e^-$  and DIS collisions.

distribution that depends on the momenta of all previous emissions. Common choices of scale for the ordering of emissions are virtuality, transverse momentum or angle. Parton showering stops at a scale of order 1 GeV, at which point a *hadronization model* is used to convert the resulting partons into hadrons. One widely-used model involves stretching a color “string” across quarks and gluons, and breaking it up into hadrons [203,204]. Another breaks each gluon into a  $q\bar{q}$  pair and then groups quarks and anti-quarks into colorless “clusters”, which then give the hadrons [198]. For  $pp$  and  $\gamma p$  processes, modeling is also needed to treat the collision between the two hadron remnants, which generates an *underlying event* (UE), usually implemented via additional  $2 \rightarrow 2$  scatterings (“multiple parton interactions”) at a scale of a few GeV, following Ref. 205.

A deficiency of the soft and collinear approximations that underlie parton showers is that they may fail to reproduce the full pattern of hard wide-angle emissions, important, for example, in many new physics searches. It is therefore common to use LO multi-parton matrix elements to generate hard high-multiplicity partonic configurations as additional starting points for the showering, supplemented with some prescription (CKKW [206], MLM [207]) for consistently merging samples with different initial multiplicities.

MCs, as described above, generate cross sections for the requested hard process that are correct at LO. A wide variety of processes are available in MC implementations that are correct to NLO, using the MC@NLO [208] or POWHEG [209] prescriptions, notably through the Madgraph5\_aMC@NLO [77], POWHEGBox [210] and Sherpa [79,211] programs. Techniques have also been developed recently to combine NLO plus shower accuracy for different multiplicities of final-state jets [212]. Building in part on some of that work, several groups have also obtained NNLO plus shower accuracy for Drell-Yan and Higgs production [213].

#### 9.2.4. Accuracy of predictions :

Estimating the accuracy of perturbative QCD predictions is not an exact science. It is often said that LO calculations are accurate to within a factor of two. This is based on experience with NLO corrections in the cases where these are available. In processes involving new partonic scattering channels at NLO and/or large ratios of scales (such as jet observables in processes with vector bosons, or the production of high- $p_t$  jets containing  $B$ -hadrons), the ratio of the NLO to LO predictions, commonly called the “ $K$ -factor”, can be substantially larger than 2.

For calculations beyond LO, a conservative approach to estimate the perturbative uncertainty is to take it to be the last known perturbative order; a more widely used method is to estimate it from the change in the prediction when varying the renormalization and factorization scales around a central value  $Q$  that is taken close to the physical scale of the process. A conventional range of variation is  $Q/2 < \mu_R, \mu_F < 2Q$ . This should not be assumed to always estimate the full uncertainty from missing higher orders, but it does indicate the size of one important known source of higher-order ambiguity.<sup>††</sup>

There does not seem to be a broad consensus on whether  $\mu_R$  and  $\mu_F$  should be kept identical or varied independently. One common option is to vary them independently with the restriction  $\frac{1}{2}\mu_R < \mu_F < 2\mu_R$  [221]. This limits the risk of misleadingly small uncertainties due to fortuitous cancellations between the  $\mu_F$  and  $\mu_R$  dependence when both are varied together, while avoiding the appearance of large logarithms of  $\mu_R^2/\mu_F^2$  when both are varied completely independently.

Calculations that involve resummations usually have an additional source of uncertainty associated with the choice of argument of the logarithms being resummed, *e.g.*  $\ln(2\frac{p_T^Z}{M_Z})$  as opposed to  $\ln(\frac{1}{2}\frac{p_T^Z}{M_Z})$ . In addition to varying renormalization and factorization scales, it is

<sup>††</sup> A number of prescriptions also exist for setting the scale automatically, *e.g.* Refs. 214–217, eliminating uncertainties from scale variation, though not from the truncation of the perturbative series itself. Recently, there have also been studies of how to estimate uncertainties from missing higher orders that go beyond scale variations [218,219,220].

therefore also advisable to vary the argument of the logarithm by a suitable factor in either direction with respect to the “natural” argument.

The accuracy of QCD predictions is limited also by non-perturbative corrections, which typically scale as a power of  $\Lambda/Q$ . For measurements that are directly sensitive to the structure of the hadronic final state, the corrections are usually linear in  $\Lambda/Q$ . The non-perturbative corrections are further enhanced in processes with a significant underlying event (*i.e.* in  $pp$  and  $p\bar{p}$  collisions) and in cases where the perturbative cross sections fall steeply as a function of  $p_t$  or some other kinematic variable, for example in inclusive jet spectra or dijet mass spectra.

Non-perturbative corrections are commonly estimated from the difference between Monte Carlo events at the parton level and after hadronization. An issue to be aware of with this procedure is that “parton level” is not a uniquely defined concept. For example, in an event generator it depends on a (somewhat arbitrary and tunable) internal cutoff scale that separates the parton showering from the hadronization. In contrast no such cutoff scale exists in a NLO or NNLO partonic calculation. For this reason there are widespread reservations as to the appropriateness of deriving hadronization corrections from a Monte Carlo program and then applying them to NLO or NNLO predictions. There exist alternative methods for estimating hadronization corrections, which attempt to analytically deduce non-perturbative effects in one observable based on measurements of other observables (see the reviews [23,222]). While they directly address the problem of different possible definitions of parton level, it should also be said that they are far less flexible than Monte Carlo programs and not always able to provide equally good descriptions of the data.

### 9.3. Experimental studies of QCD

Since we are not able to directly measure partons (quarks or gluons), but only hadrons and their decay products, a central issue for every experimental study of perturbative QCD is establishing a correspondence between observables obtained at the partonic and the hadronic level. The only theoretically sound correspondence is achieved by means of *infrared and collinear safe* quantities, which allow one to obtain finite predictions at any order of perturbative QCD.

As stated above, the simplest case of infrared- and collinear-safe observables are total cross sections. More generally, when measuring fully inclusive observables, the final state is not analyzed at all regarding its (topological, kinematical) structure or its composition. Basically the relevant information consists in the rate of a process ending up in a partonic or hadronic final state. In  $e^+e^-$  annihilation, widely used examples are the ratios of partial widths or branching ratios for the electroweak decay of particles into hadrons or leptons, such as  $Z$  or  $\tau$  decays, (*cf.* Sec. 9.2.1). Such ratios are often favored over absolute cross sections or partial widths because of large cancellations of experimental and theoretical systematic uncertainties. The strong suppression of non-perturbative effects,  $\mathcal{O}(\Lambda^4/Q^4)$ , is one of the attractive features of such observables, however, at the same time the sensitivity to radiative QCD corrections is small, which for example affects the statistical uncertainty when using them for the determination of the strong coupling constant. In the case of  $\tau$  decays not only the hadronic branching ratio is of interest, but also moments of the spectral functions of hadronic tau decays, which sample different parts of the decay spectrum and thus provide additional information. Other examples of fully inclusive observables are structure functions (and related sum rules) in DIS. These are extensively discussed in Sec. 19 of this *Review*.

On the other hand, often the structure or composition of the final state are analyzed and cross sections differential in one or more variables characterizing this structure are of interest. Examples are jet rates, jet substructure, event shapes or transverse momentum distributions of jets or vector bosons in hadron collisions. The case of fragmentation functions, *i.e.* the measurement of hadron production as a function of the hadron momentum relative to some hard scattering scale, is discussed in Sec. 20 of this *Review*.

It is worth mentioning that, besides the correspondence between the parton and hadron level, also a correspondence between the hadron level and the actually measured quantities in the detector has to be established. The simplest examples are corrections for finite experimental acceptance and efficiencies. Whereas acceptance corrections essentially are of theoretical nature, since they involve extrapolations from the measurable (partial) to the full phase space, other corrections such as for efficiency, resolution and response, are of experimental nature. For example, measurements of differential cross sections such as jet rates require corrections in order to relate, *e.g.* the energy deposits in a calorimeter to the jets at the hadron level. Typically detector simulations and/or data-driven methods are used in order to obtain these corrections. Care should be taken here in order to have a clear separation between the parton-to-hadron level and hadron-to-detector level corrections. Finally, for the sake of an easy comparison to the results of other experiments and/or theoretical calculations, it is suggested to provide, whenever possible, measurements corrected for detector effects and/or all necessary information related to the detector response (*e.g.* the detector response matrix).

### 9.3.1. Hadronic final-state observables :

#### 9.3.1.1. Jets:

In hard interactions, final-state partons and hadrons appear predominantly in collimated bunches, which are generically called *jets*. To a first approximation, a jet can be thought of as a hard parton that has undergone soft and collinear showering and then hadronization. Jets are used both for testing our understanding and predictions of high-energy QCD processes, and also for identifying the hard partonic structure of decays of massive particles like top quarks.

In order to map observed hadrons onto a set of jets, one uses a *jet definition*. The mapping involves explicit choices: for example when a gluon is radiated from a quark, for what range of kinematics should the gluon be part of the quark jet, or instead form a separate jet? Good jet definitions are infrared and collinear safe, simple to use in theoretical and experimental contexts, applicable to any type of inputs (parton or hadron momenta, charged particle tracks, and/or energy deposits in the detectors) and lead to jets that are not too sensitive to non-perturbative effects.

An extensive treatment of the topic of jet definitions is given in Ref. 223 (for  $e^+e^-$  collisions) and Refs. [224–226]. Here we briefly review the two main classes: cone algorithms, extensively used at older hadron colliders, and sequential recombination algorithms, more widespread in  $e^+e^-$  and  $ep$  colliders and at the LHC.

Very generically, most (iterative) cone algorithms start with some seed particle  $i$ , sum the momenta of all particles  $j$  within a cone of opening-angle  $R$ , typically defined in terms of (pseudo-)rapidity and azimuthal angle. They then take the direction of this sum as a new seed and repeat until the cone is stable, and call the contents of the resulting stable cone a jet if its transverse momentum is above some threshold  $p_{t,\min}$ . The parameters  $R$  and  $p_{t,\min}$  should be chosen according to the needs of a given analysis.

There are many variants of cone algorithm, and they differ in the set of seeds they use and the manner in which they ensure a one-to-one mapping of particles to jets, given that two stable cones may share particles (“overlap”). The use of seed particles is a problem w.r.t. infrared and collinear safety, and seeded algorithms are generally not compatible with higher-order (or sometimes even leading-order) QCD calculations, especially in multi-jet contexts, as well as potentially subject to large non-perturbative corrections and instabilities. Seeded algorithms (JetCLU, MidPoint, and various other experiment-specific iterative cone algorithms) are therefore to be deprecated. A modern alternative is to use a seedless variant, SIScone [227].

Sequential recombination algorithms at hadron colliders (and in DIS) are characterized by a distance  $d_{ij} = \min(k_{t,i}^{2p}, k_{t,j}^{2p})\Delta_{ij}^2/R^2$  between all pairs of particles  $i, j$ , where  $\Delta_{ij}$  is their separation in the rapidity-azimuthal plane,  $k_{t,i}$  is the transverse momentum

w.r.t. the incoming beams, and  $R$  is a free parameter. They also involve a “beam” distance  $d_{iB} = k_{t,i}^{2p}$ . One identifies the smallest of all the  $d_{ij}$  and  $d_{iB}$ , and if it is a  $d_{ij}$ , then  $i$  and  $j$  are merged into a new pseudo-particle (with some prescription, a recombination scheme, for the definition of the merged four-momentum). If the smallest distance is a  $d_{iB}$ , then  $i$  is removed from the list of particles and called a jet. As with cone algorithms, one usually considers only jets above some transverse-momentum threshold  $p_{t,\min}$ . The parameter  $p$  determines the kind of algorithm:  $p = 1$  corresponds to the (*inclusive*-) $k_t$  algorithm [154,228,229],  $p = 0$  defines the *Cambridge-Aachen* algorithm [230,231], while for  $p = -1$  we have the *anti- $k_t$*  algorithm [232]. All these variants are infrared and collinear safe to all orders of perturbation theory. Whereas the former two lead to irregularly shaped jet boundaries, the latter results in cone-like boundaries. The *anti- $k_t$*  algorithm has become the de-facto standard for the LHC experiments.

In  $e^+e^-$  annihilations the  $k_t$  algorithm [154] uses  $y_{ij} = 2 \min(E_i^2, E_j^2)(1 - \cos\theta_{ij})/Q^2$  as distance measure and repeatedly merges the pair with smallest  $y_{ij}$ , until all  $y_{ij}$  distances are above some threshold  $y_{\text{cut}}$ , the jet resolution parameter. The (pseudo)-particles that remain at this point are called the jets. Here it is  $y_{\text{cut}}$  (rather than  $R$  and  $p_{t,\min}$ ) that should be chosen according to the needs of the analysis. As mentioned earlier, the  $k_t$  algorithm has the property that logarithms  $\ln(1/y_{\text{cut}})$  exponentiate in resummation calculations. This is one reason why it is preferred over the earlier JADE algorithm [156], which uses the distance measure  $y_{ij} = 2 E_i E_j (1 - \cos\theta_{ij})/Q^2$ . Note that other variants of sequential recombination algorithms for  $e^+e^-$  annihilations, using different definitions of the resolution measure  $y_{ij}$ , exhibit much larger sensitivities to fragmentation and hadronization effects than the  $k_t$  and JADE algorithms [233].

Efficient implementations of the above algorithms are available through the *FastJet* package [234].

#### 9.3.1.2. Event Shapes:

Event-shape variables are functions of the four momenta of the particles in the final state and characterize the topology of an event’s energy flow. They are sensitive to QCD radiation (and correspondingly to the strong coupling) insofar as gluon emission changes the shape of the energy flow.

The classic example of an event shape is the *Thrust* [235,236] in  $e^+e^-$  annihilations, defined as

$$\hat{\tau} = \max_{\vec{n}_\tau} \frac{\sum_i |\vec{p}_i \cdot \vec{n}_\tau|}{\sum_i |\vec{p}_i|}, \quad (9.22)$$

where  $\vec{p}_i$  are the momenta of the particles or the jets in the final-state and the maximum is obtained for the Thrust axis  $\vec{n}_\tau$ . In the Born limit of the production of a perfect back-to-back  $q\bar{q}$  pair the limit  $\hat{\tau} \rightarrow 1$  is obtained, whereas a perfectly spherical many-particle configuration leads to  $\hat{\tau} \rightarrow 1/2$ . Further event shapes of similar nature have been extensively measured at LEP and at HERA, and for their definitions and reviews we refer to Refs. 1,4,222,237,238. The energy-energy correlation function [239], namely the energy-weighted angular distribution of produced hadron pairs, and its associated asymmetry are further shape variables that have been studied in detail at  $e^+e^-$  colliders. For hadron colliders the appropriate modification consists in only taking the transverse momentum component [240]. More recently, the event shape *N-jettiness* has been proposed [241], that measures the degree to which the hadrons in the final state are aligned along  $N$  jet axes or the beam direction. It vanishes in the limit of exactly  $N$  infinitely narrow jets.

Phenomenological discussions of event shapes at hadron colliders can be found in Refs. [241–243]. Measurements of hadronic event-shape distributions have been published by CDF [244], ATLAS [245–247] and CMS [248–250].

Event shapes are used for many purposes. These include measuring the strong coupling, tuning the parameters of Monte Carlo programs, investigating analytical models of hadronization and distinguishing QCD events from events that might involve decays of new particles (giving event-shape values closer to the spherical limit).

### 9.3.1.3. Jet substructure, quark vs. gluon jets:

Jet substructure, which can be resolved by finding subjets or by measuring jet shapes, is sensitive to the details of QCD radiation in the shower development inside a jet and has been extensively used to study differences in the properties of quark and gluon induced jets, strongly related to their different color charges. In general there is clear experimental evidence that gluon jets have a softer particle spectrum and are “broader” than (light-) quark jets, when looking at observables such as the jet shape  $\Psi(r/R)$ . This is the fractional transverse momentum contained within a sub-cone of cone-size  $r$  for jets of cone-size  $R$ . It is sensitive to the relative fractions of quark and gluon jets in an inclusive jet sample and receives contributions from soft-gluon initial-state radiation and the underlying event. Therefore, it has been widely employed for validation and tuning of Monte Carlo models. Furthermore, this quantity turns out to be sensitive to the modification of the gluon radiation pattern in heavy ion collisions (see e.g. Ref. 251).

The most recent jet shape measurements using proton-proton collision data have been presented for inclusive jet samples [252,253] and for top-quark production [254]. Further discussions, references and recent summaries can be found in Refs. 238, 255, 256 and Sec. 4 of Ref. 257.

The use of jet substructure has also been investigated in order to distinguish QCD jets from jets that originate from hadronic decays of boosted massive particles (high- $p_t$  electroweak bosons, top quarks and hypothesized new particles). Recently, a considerable number of experimental studies have been carried out with Tevatron and LHC data, in order to investigate on the performance of the proposed algorithms for resolving jet substructure and to apply them to searches for new physics. For reviews of this rapidly growing field, see sec. 5.3 of Ref. 224, Ref. 226 and Refs. [257–260].

### 9.3.2. QCD measurements at colliders :

There exists a wealth of data on QCD-related measurements in  $e^+e^-$ ,  $ep$ ,  $pp$ , and  $p\bar{p}$  collisions, to which a short overview like this would not be able to do any justice. Extensive reviews of the subject have been published in Refs. 237, 238 for  $e^+e^-$  colliders and in Ref. 261 for  $ep$  scattering, whereas for hadron colliders comprehensive overviews are given in, e.g., Refs. 225, 256 and Refs. [262–264].

Below we concentrate our discussion on measurements that are most sensitive to hard QCD processes, with focus on jet production.

**9.3.2.1.  $e^+e^-$  colliders:** Analyses of jet production in  $e^+e^-$  collisions are mostly based on data from the JADE experiment at center-of-mass energies between 14 and 44 GeV, as well as on LEP collider data at the  $Z$  resonance and up to 209 GeV. They cover the measurements of (differential or exclusive) jet rates (with multiplicities typically up to 4, 5 or 6 jets), the study of 3-jet events and particle production between the jets as a tool for testing hadronization models, as well as 4-jet production and angular correlations in 4-jet events.

Event-shape distributions from  $e^+e^-$  data have been an important input to the tuning of parton shower MC models, typically matched to matrix elements for 3-jet production. In general these models provide good descriptions of the available, highly precise data. Especially for the large LEP data sample at the  $Z$  peak, the statistical uncertainties are mostly negligible and the experimental systematic uncertainties are at the percent level or even below. These are usually dominated by the uncertainties related to the MC model dependence of the efficiency and acceptance corrections (often referred to as “detector corrections”).

Observables measured in  $e^+e^-$  collisions have been used for determinations of the strong coupling constant (*cf.* Section 9.4 below) and for putting constraints on the QCD color factors (*cf.* Sec. 9.1 for their definitions), thus probing the non-abelian nature of QCD. Typically, cross sections can be expressed as functions of these color factors, for example  $\sigma = f(\alpha_s C_F, C_A/C_F, n_f T_R/C_F)$ . Angular correlations in 4-jet events give sensitivity at leading order. Some sensitivity to these color factors, although only at NLO, is also obtained from event-shape distributions. Scaling violations of fragmentation functions and the different subjet structure in quark and gluon induced jets also give access to these color factors. In order to

extract absolute values, e.g. for  $C_F$  and  $C_A$ , certain assumptions have to be made for other parameters, such as  $T_R, n_f$  or  $\alpha_s$ , since typically only combinations (ratios, products) of all the relevant parameters appear in the perturbative predictions. A compilation of results [238] quotes world average values of  $C_A = 2.89 \pm 0.03(\text{stat}) \pm 0.21(\text{syst})$  and  $C_F = 1.30 \pm 0.01(\text{stat}) \pm 0.09(\text{syst})$ , with a correlation coefficient of 82%. These results are in perfect agreement with the expectations from SU(3) of  $C_A = 3$  and  $C_F = 4/3$ .

**9.3.2.2. DIS and photoproduction:** Multi-jet production in  $ep$  collisions at HERA, both in the DIS and photoproduction regime, allows for tests of QCD factorization (one initial-state proton and its associated PDF versus the hard scattering which leads to high- $p_t$  jets) and NLO calculations which exist for 2- and 3-jet final states. Sensitivity is also obtained to the product of the coupling constant and the gluon PDF. Experimental uncertainties of the order of 5–10% have been achieved, mostly dominated by the jet energy scale, whereas statistical uncertainties are negligible to a large extent. For comparison to theoretical predictions, at large jet  $p_t$  the PDF uncertainty dominates the theoretical uncertainty (typically of order 5–10%, in some regions of phase space up to 20%), therefore jet observables become useful inputs for PDF fits.

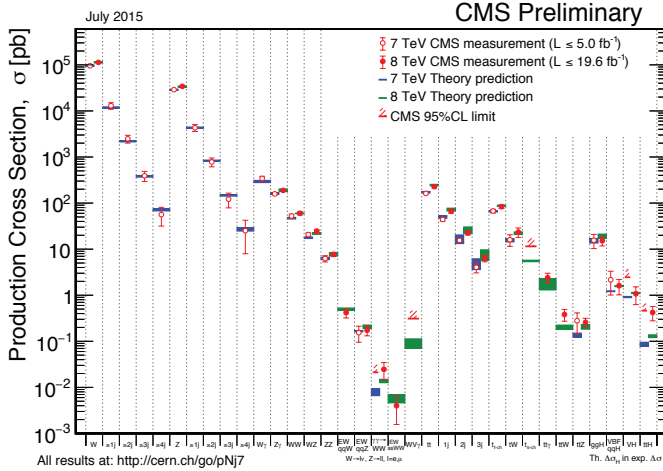
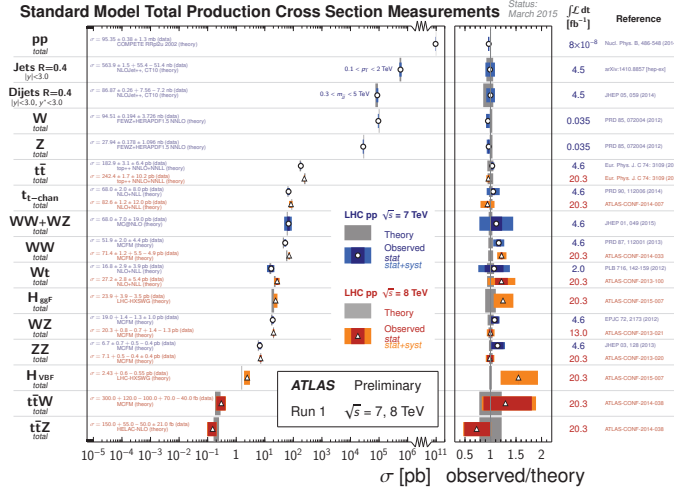
In general, the data are well described by NLO matrix-element calculations, combined with DGLAP evolution equations, in particular at large  $Q^2$  and central values of jet pseudo-rapidity. At low values of  $Q^2$  and  $x$ , in particular for large jet pseudo-rapidities, certain features of the data have been interpreted as requiring BFKL-type evolution, though the predictions for such schemes are still limited. It is worth noting that there is lack of consensus throughout the community regarding this need of BFKL-evolution at currently probed  $x, Q^2$  values, and an alternative approach [265] that implements the merging of LO matrix-element based event generation with a parton shower (using the SHERPA framework) successfully describes the data in all kinematical regions, including the low  $Q^2$ , low  $x$  domain. At moderately small  $x$  values, it should perhaps not be surprising that the BFKL approach and fixed-order matrix-element merging with parton showers may both provide adequate descriptions of the data, because some part of the multi-parton phase space that they model is common to both approaches.

In the case of photoproduction, a wealth of measurements with low  $p_t$  jets were performed in order to constrain the photon PDFs. The uncertainties related to these photon PDFs play a minor role at high jet  $p_t$ , which has allowed for precise tests of pQCD calculations.

A few examples of recent measurements can be found in Refs. 266–274 for DIS and in Refs. 275–279 for photoproduction.

**9.3.2.3. Hadron colliders:** The spectrum of observables and the number of measurements performed at hadron colliders is enormous, probing many regions of phase space and covering a huge range of cross sections, as illustrated in Fig. 9.1 for the case of the ATLAS and CMS experiments at the LHC. For the sake of brevity, in the following only certain classes of those measurements will be discussed, that allow addressing particular aspects of the various QCD studies performed. Most of our discussion will focus on recent LHC results, which are available for center-of-mass energies of 2.76, 7 and 8 TeV with integrated luminosities of up to  $20 \text{ fb}^{-1}$ . Generally speaking, besides representing a general test of the standard model and QCD in particular, these measurements serve several purposes, such as: (i) probing pQCD and its various approximations and implementations in MC models, in order to quantify the order of magnitude of not yet calculated contributions and to gauge their precision when used as background predictions, or (ii) extracting/constraining model parameters such as the strong coupling constant or PDFs.

Among the most important cross sections measured is the inclusive jet spectrum as a function of the jet transverse energy ( $E_t$ ) or the jet transverse momentum ( $p_t$ ), for several rapidity regions and for  $p_t$  up to 700 GeV at the Tevatron and  $\sim 2$  TeV at the LHC. It is worth noting that this upper limit in  $p_t$  corresponds to a distance scale of  $\sim 10^{-19}$  m: no other experiment so far is able to directly probe smaller distance scales of nature than this measurement. Whereas the Tevatron measurements (Refs. 282–284) were based on the infrared-



**Figure 9.1:** Overview of cross section measurements for a wide class of processes and observables, as obtained by the ATLAS [280] and CMS [281] experiments at the LHC, for centre-of-mass energies of 7 and 8 TeV. Also shown are the theoretical predictions and their uncertainties.

and collinear-safe  $k_t$  algorithm in addition to the more widely used Midpoint and JetCLU algorithms of the past, the LHC experiments focus on the *anti- $k_t$*  algorithm using various radius parameters. Recent measurements by ALICE, ATLAS and CMS have been published in Refs. 285–288. Reviews can be found in, *e.g.*, Refs. 256,289,290 and a compilation of inclusive jet cross section results is presented in Fig. 51.1 [291] of this Review.

In general we observe a good description of the data by the NLO QCD predictions, over about 11 orders of magnitude in cross section. The experimental systematic uncertainties are dominated by the jet energy scale uncertainty, quoted to be in the range of a few percent (see for instance the review in Ref. 292), leading to uncertainties of  $\sim 5 - 30\%$  on the cross section, increasing with  $p_t$  and rapidity. The PDF uncertainties dominate the theoretical uncertainty at large  $p_t$  and rapidity. In fact, inclusive jet data are important inputs to global PDF fits (see [293] for a recent review). Constraints on the PDFs can also be obtained from ratios of inclusive cross sections at different center-of-mass energies [286]. In general, ratios of jet cross sections are a means to (at least partially) cancel the jet energy scale uncertainties and thus provide jet observables with significantly improved precision.

Dijet events are analyzed in terms of their invariant mass and angular distributions, which allows for tests of NLO QCD predictions (see *e.g.* Refs. [288,294] for recent LHC results), and setting stringent limits on deviations from the Standard Model, such as quark

compositeness (some examples can be found in Refs. 295–299). Furthermore, dijet azimuthal correlations between the two leading jets, normalized to the total dijet cross section, are an extremely valuable tool for studying the spectrum of gluon radiation in the event. The azimuthal separation of the two leading jets is sensitive to multi-jet production, avoiding at the same time large systematic uncertainties from the jet energy calibration. For example, results from the Tevatron [300,301] and the LHC [302,303] show that the LO (non-trivial) prediction for this observable, with at most three partons in the final state, is not able to describe the data for an azimuthal separation below  $2\pi/3$ , where NLO contributions (with 4 partons) restore the agreement with data. In addition, this observable can be employed to tune Monte Carlo predictions of soft gluon radiation.

Further examples of dijet observables that probe special corners of phase space are those which involve forward (large rapidity) jets and where a large rapidity separation, possibly also a rapidity gap, is required between the two jets. Reviews of such measurements can be found in Refs. [256,304], showing that no single prediction is capable of describing the data in all phase-space regions. In particular, no conclusive evidence for BFKL effects in these observables has been established so far.

Beyond dijet final states, measurements of the production of three or more jets, including cross section ratios, have been performed (see Refs. [256,305] for recent reviews), as a means of testing perturbative QCD predictions, determining the strong coupling constant (at NLO precision so far), and probing/tuning MC models, in particular those combining multi-parton matrix elements with parton showers.

In terms of precision achieved, measurements of inclusive vector boson ( $W, Z$ ) production outperform the jet studies described above and provide the most precisely determined observables at hadron colliders so far. This is because the experimental signatures are based on leptons that are measured much more accurately than jets. At the LHC [306–310], the dominant uncertainty stems from the luminosity determination ( $\sim 2-4\%$ ), while other uncertainties (*e.g.* statistics, lepton efficiencies) are controlled at the 1–3% level. The uncertainty from the acceptance correction of about 1–2% can be reduced by measuring so-called fiducial cross sections, *ie.* by applying kinematic cuts also to the particle level of the theoretical predictions. A further reduction or even complete elimination of particular uncertainties (*e.g.* luminosity) is achieved by measuring cross section ratios ( $W/Z$  or  $W^+/W^-$ ) or differential distributions that are normalised to the inclusive cross section.

On the theory side, as discussed earlier in this review, the production of these color-singlet states has been calculated up to NNLO accuracy. Since currently the dominant theoretical uncertainty (of order 5% and thus similar or larger than the experimental errors) is related to the choice of PDFs, these data provide useful handles for PDF determinations.

Further insights are obtained from measurements of differential vector boson production, as a function of the invariant dilepton mass, the boson's rapidity or its transverse momentum. For example, the dilepton invariant mass distribution has been measured [311–314] for masses between 15 and 2000 GeV, covering more than 8 orders of magnitude in cross section. NNLO QCD predictions, together with modern PDF sets and including higher-order electroweak and QED final-state radiation corrections, describe the data to within 5–10% over this large range, whereas NLO predictions show larger deviations, unless matched to a parton shower.

Similar conclusions can be drawn from the observed rapidity distribution of the dilepton system or, in the case of  $W$  production, from the observed charged lepton rapidity distribution and its charge asymmetry. The latter is particularly sensitive to differences among PDF sets [315], also thanks to the high precision achieved by the ATLAS and CMS experiments for central rapidity ranges. These measurements are nicely extended to the very forward region, up to 4.5 in lepton rapidity, by the LHCb experiment.

An overview of recent results can be found in Ref. 256. There one can also find a discussion of and references to recent LHC results from studies of the vector boson's transverse momentum distribution,  $p_t^V$ . This observable probes different aspects of higher-order QCD

effects and is sensitive to jet production in association to the vector boson, without suffering from the large jet energy scale uncertainties since there is no explicit jet reconstruction. Whereas in the  $p_t^V$  region of several tens to hundreds of GeV the NNLO predictions (that effectively are of NLO accuracy for this variable) agree with the data to within about 10%, at transverse momentum below  $\sim 5$ –10 GeV the fixed order predictions fail and soft-gluon resummation is needed to restore the agreement with data. Correspondingly, MC models implementing parton shower matching to LO or NLO matrix elements provide good predictions at low and intermediate  $p_t^V$ , but deviate up to 40% at high  $p_t^V$ .

While in principle inclusive and differential photon production represents a similar tool for studying effects as described above, the experimental results are less precise than for  $W$  and  $Z$  production, related to the greater challenges encountered in photon reconstruction and purity determination compared to lepton final states.

In terms of complexity, probably the most challenging class of processes is vector boson (photon,  $W$ ,  $Z$ ) production together with jets. By now the amount of results obtained both at the Tevatron and at the LHC is so extensive that a comprehensive discussion with a complete citation list would go much beyond the scope of this *Review*. We rather refer to recent summaries such as those in Refs. 256,316 and to previous versions of this *Review*.

The measurements cover a very large phase space, *e.g.* with jet transverse momenta between 30 GeV and  $\sim 800$  GeV and jet rapidities up to  $|y| < 4.4$ . Jet multiplicities as high as seven jets accompanying the vector boson have already been probed at the LHC, together with a substantial number of other kinematical observables, such as angular correlations among the various jets or among the jets and the vector boson, or the sum of jet transverse momenta,  $H_T$ . Whereas the jet  $p_t$  and  $H_T$  distributions are dominated by jet energy scale uncertainties at levels similar to those discussed above for inclusive jet production, angular correlations and jet multiplicity ratios have been measured with a precision of  $\sim 10\%$ , see *e.g.* Refs. 317,249.

A general observation is that MC models, which implement a matching of matrix-element calculations with parton showers, provide a good description of the data within uncertainties. Also NLO calculations for up to five jets [110] in addition to the vector boson are in good agreement with the data over that phase space, where the calculations are applicable; that is, one can not expect such predictions to work for, *e.g.*, the  $p_t$  distribution of the  $n+1$ st jet with  $V+n$  jets calculated at NLO. However, with the high statistics available to and the high precision achieved by the LHC experiments, some more detailed observations can be made. MC models that implement parton shower matching to LO matrix elements (LO+PS) tend to overpredict the data at large jet and/or boson  $p_t$ , while parton shower matching to NLO matrix elements gives better agreement. These problems of LO+PS models are less acute when looking at angular correlations.

Also, electroweak corrections are expected to become more and more relevant now that the TeV energy range starts to be explored. For example, such corrections were found [318] to be sizeable (tens of percent) when studying the ratio  $(d\sigma^\gamma/dp_t)/(d\sigma^Z/dp_t)$  in  $\gamma(Z)+\text{jet}$  production,  $p_t$  being the boson's transverse momentum, and might account for (some of) the differences observed in a recent CMS measurement [319] of this quantity.

The challenges get even more severe in the case of vector boson plus heavy quark ( $b$ ,  $c$ ) production, both because of theoretical issues (an additional scale is introduced by the heavy quark mass and different schemes exist for the handling of heavy quarks and their mass effects in the initial and/or final state) and because of additional experimental uncertainties related to the heavy-flavour tagging. A recent review of heavy quark production at the LHC can be found in Ref. 320. There it is stated that studies of  $b$ -jet production with or without associated  $W$  and  $Z$  bosons reveal the di- $b$ -jet  $p_t$  and mass spectra to be well modelled, within experimental and theoretical uncertainties, by most generators on the market. However, sizeable differences between data and predictions are seen in the modelling of events with single  $b$  jets, particularly at large  $b$ -jet  $p_t$ , where gluon splitting processes become dominant, as also confirmed by studies of  $b$ -hadron and  $b$ -jet angular correlations.

A number of interesting developments, in terms of probing higher-order QCD effects, have recently occurred in the sector of diboson production, in particular for the  $WW$  and  $\gamma\gamma$  cases. Regarding the former, a disagreement of about 10% between the LHC measurements and the NLO predictions has led to a number of speculations of possible new physics effects in this channel. However, the latest CMS measurement [321], with a relative accuracy of 8%, is in excellent agreement with a recent NNLO calculation [56], that has an estimated 3% residual uncertainty from missing contributions beyond NNLO.

In the case of diphoton production, ATLAS [322] and CMS [323] have provided accurate measurements, in particular for phase-space regions that are sensitive to radiative QCD corrections (multi-jet production), such as small azimuthal photon separation. While there are large deviations between data and NLO predictions in this region, a calculation [132] at NNLO accuracy manages to mostly fill this gap. This is an interesting example where scale variations can not provide a reliable estimate of missing contributions beyond NLO, since at NNLO new channels appear in the initial state (gluon fusion in this case).

In terms of heaviest particle involved, top-quark production at the LHC has become an important tool for probing higher-order QCD calculations, thanks to very impressive achievements both on the experimental and theoretical side, as extensively summarised in Ref. 324. Regarding  $t\bar{t}$  production, the most precise inclusive cross section measurements are achieved using the dilepton ( $e\mu$ ) final state, with a total uncertainty of 4%. This is of about the same size as the uncertainty on the most advanced theoretical prediction [55], obtained at NNLO with additional soft-gluon resummation at NNLL accuracy. There is excellent agreement between data and QCD prediction.

A large number of differential cross section measurements have been performed at 7 and 8 TeV centre-of-mass energy, studying distributions such as the top-quark  $p_t$  and rapidity, the transverse momentum and invariant mass of the  $t\bar{t}$  system (probing scales up to the TeV range), or the number of additional jets. These measurements have been compared to a wide range of predictions, at fixed order up to NNLO as well as using LO or NLO matrix elements matched to parton showers. While in general there is good agreement observed with data, most MC simulations predict a somewhat harder top-quark  $p_t$  distribution than seen in data, an effect that is currently under investigation.

Thanks to both the precise measurements of and predictions for the inclusive top-pair cross section, that is sensitive to the strong coupling constant and the top-quark mass, this observable has been used to measure the strong coupling constant for the first time at NNLO accuracy from hadron collider data [325] (*cf.* Section 9.4 below), as well as to obtain a measurement of the top-quark's pole mass without employing direct reconstruction methods [325,326].

Finally, it is worth mentioning that first steps are being undertaken towards using the newly found Higgs boson as a new tool for QCD studies, since Higgs production, dominated by the gluon fusion process, is subject to very large QCD corrections. First studies of fiducial and differential cross sections, using the  $ZZ$  and  $\gamma\gamma$  decay channels, have already been performed [327–329], and the current experimental precision of  $\sim 20\%$  or more is expected to be substantially reduced with the future LHC data.

## 9.4. Determinations of the strong coupling constant

Beside the quark masses, the only free parameter in the QCD Lagrangian is the strong coupling constant  $\alpha_s$ . The coupling constant in itself is not a physical observable, but rather a quantity defined in the context of perturbation theory, which enters predictions for experimentally measurable observables, such as  $R$  in Eq. (9.7).

Many experimental observables are used to determine  $\alpha_s$ . Considerations in such determinations include:

- The observable's sensitivity to  $\alpha_s$  as compared to the experimental precision. For example, for the  $e^+e^-$  cross section to hadrons (*cf.*  $R$  in Sec. 9.2.1), QCD effects are only a small correction, since the perturbative series starts at order  $\alpha_s^0$ ; 3-jet production or event shapes in  $e^+e^-$  annihilations are directly sensitive to

$\alpha_s$  since they start at order  $\alpha_s$ ; the hadronic decay width of heavy quarkonia,  $\Gamma(\Upsilon \rightarrow \text{hadrons})$ , is very sensitive to  $\alpha_s$  since its leading order term is  $\propto \alpha_s^3$ .

- The accuracy of the perturbative prediction, or equivalently of the relation between  $\alpha_s$  and the value of the observable. The minimal requirement is generally considered to be an NLO prediction. Some observables are predicted to NNLO (many inclusive observables, 3-jet rates and event shapes in  $e^+e^-$  collisions) or even N<sup>3</sup>LO ( $e^+e^-$  hadronic cross section and  $\tau$  branching fraction to hadrons). In certain cases, fixed-order predictions are supplemented with resummation. The precise magnitude of theory uncertainties is usually estimated as discussed in Sec. 9.2.4.
- The size of non-perturbative effects. Sufficiently inclusive quantities, like the  $e^+e^-$  cross section to hadrons, have small non-perturbative contributions  $\sim \Lambda^4/Q^4$ . Others, such as event-shape distributions, have contributions  $\sim \Lambda/Q$ .
- The scale at which the measurement is performed. An uncertainty  $\delta$  on a measurement of  $\alpha_s(Q^2)$ , at a scale  $Q$ , translates to an uncertainty  $\delta' = (\alpha_s^2(M_Z^2)/\alpha_s^2(Q^2)) \cdot \delta$  on  $\alpha_s(M_Z^2)$ . For example, this enhances the already important impact of precise low- $Q$  measurements, such as from  $\tau$  decays, in combinations performed at the  $M_Z$  scale.

In this review, we update the measurements of  $\alpha_s$  summarized in the 2013 edition, and we extract a new world average value of  $\alpha_s(M_Z^2)$  from the most significant and complete results available today<sup>‡</sup>.

We restrict the selection of results from which to determine the world average value of  $\alpha_s(M_Z^2)$  to those which are

- published in a peer-reviewed journal,
- based on the most complete perturbative QCD predictions, *i.e.* to those using NNLO or higher-order expansions.

This excludes *e.g.* results from jet production in DIS at HERA and at hadron colliders, for which calculations are available at NLO only. These will nevertheless be discussed in this review, as they are important ingredients for the experimental evidence of the energy dependence of  $\alpha_s$ , *i.e.* for Asymptotic Freedom, one of the key features of QCD. Note that results which do not include reliable estimates of experimental, systematic and theoretical uncertainties, which are based on not commonly accepted procedures like scale optimization, or which omit discussion or accounting of non-perturbative corrections and effects, will not be referenced at all in this review.

In order to calculate the world average value of  $\alpha_s(M_Z^2)$ , we apply an intermediate step of pre-averaging results within certain sub-fields like  $e^+e^-$  annihilation, DIS and hadronic  $\tau$ -decays, and calculate the overall world average from those pre-averages rather than from individual measurements. This is done because in most sub-fields one observes that different determinations of the strong coupling from substantially similar datasets lead to values of  $\alpha_s$  that are only marginally compatible with each other, or with the final world average value, which presumably is a reflection of the challenges of evaluating and including appropriate systematic uncertainties.

So for each sub-field, the *unweighted average* of all selected results is taken as the pre-average value of  $\alpha_s(M_Z^2)$ , and the unweighted average of the quoted uncertainties is assigned to be the respective overall error of this pre-average. However, if this error appears to be smaller than the unweighted standard deviation - *i.e.* the *spread* - of the results, the standard deviation is taken as the overall uncertainty instead. This is done in order to arrive at an unbiased estimator of the average value of  $\alpha_s(M_Z^2)$  from this sub-field, and to avoid that singular, optimistic estimates of systematic uncertainties dominate the field if these are not backed up by a broader consensus<sup>†</sup>.

Assuming that the resulting pre-averages are largely independent of each other, we determine the final world average value using the method of ‘ $\chi^2$  averaging’, as proposed, *e.g.*, in Ref. 333, in order

<sup>‡</sup> The time evolution of  $\alpha_s$  combinations can be followed by consulting Refs. [330–332] as well as earlier editions of this *Review*.

<sup>†</sup> In most practical cases, this procedure arrives at similar values as obtained from the ‘*range averaging*’ method which we used in previous *Reviews*, while it avoids potential shortcomings and biases of the latter.

to treat cases of possible (unknown) correlations as well as possibly underestimated systematic uncertainties in a meaningful and well defined manner: the central value is determined as the weighted average of the different input values. An initial uncertainty of the central value is determined treating the uncertainties of all individual measurements as being uncorrelated and of Gaussian nature, and the overall  $\chi^2$  to the central value is calculated. If this initial  $\chi^2$  is larger than the number of degrees of freedom, then all individual uncertainties are enlarged by a common factor such that  $\chi^2/\text{d.o.f.}$  equals unity. If the initial value of  $\chi^2$  is smaller than the number of degrees of freedom, an overall correlation coefficient is introduced and determined by requiring that the total  $\chi^2/\text{d.o.f.}$  equals unity. In both cases, the resulting overall uncertainty of  $\alpha_s$  is larger than the initial estimate of the uncertainty.

#### 9.4.1. Hadronic $\tau$ decays :

Based on complete N<sup>3</sup>LO predictions [28], analyses of the  $\tau$  hadronic decay width and spectral functions have been performed, leading to precise determinations of  $\alpha_s$  at the energy scale of  $M_\tau^2$  [28,334–340]. They are based on different approaches to treat perturbative and non-perturbative contributions, the impacts of which are a matter of intense discussions, see *e.g.* [338] and [341].

In particular, there is a significant difference between results obtained using fixed-order (FOPT) or contour improved perturbation theory (CIPT), such that analyses based on CIPT generally arrive at about 7% larger values of  $\alpha_s(M_\tau^2)$  than those based on FOPT. When converted to  $\alpha_s(M_Z^2)$ , the difference is about 2%. This uncertainty is about 5 times larger than the typically achieved experimental precision. In addition, most recent results show differences of up to 10% in  $\alpha_s(M_\tau^2)$  (3% at  $M_Z$ ), between different groups using the same data sets and perturbative calculations, most likely due to different treatments of the non-perturbative contributions, *c.f.* Ref. [340] with Refs. [338,339].

We determine the pre-average value of  $\alpha_s(M_Z^2)$  for this sub-field from studies which employ both, FOPT and CIPT expansions, and which include the difference among these in the quoted overall uncertainty:  $\alpha_s(M_Z^2) = 0.1202 \pm 0.0019$  [28],  $\alpha_s(M_Z^2) = 0.1200 \pm 0.0015$  [338],  $\alpha_s(M_Z^2) = 0.1199 \pm 0.0015$  [339], and  $\alpha_s(M_Z^2) = 0.1165 \pm 0.0019$  [340].

We also include the result from  $\tau$  decay and lifetime measurements, obtained in Sec. *Electroweak Model and constraints on New Physics* of the 2013 edition of this *Review*,  $\alpha_s(M_Z^2) = 0.1193 \pm 0.0023$ . All these are summarised in Fig. 9.2. Determining the unweighted average of the central values and their overall uncertainties, we arrive at  $\alpha_s(M_Z^2) = 0.1192 \pm 0.0018$  which we will use as the first input for determining the world average value of  $\alpha_s(M_Z^2)$ . This corresponds to  $\alpha_s(M_\tau^2) = 0.325 \pm 0.015$  at the scale of the  $\tau$ -mass.

#### 9.4.2. Lattice QCD :

There are several current results on  $\alpha_s$  from lattice QCD, see also Sec. *Lattice QCD* in this *Review*. The HPQCD collaboration [342] computes Wilson loops and similar short-distance quantities with lattice QCD and analyzes them with NNLO perturbative QCD. This yields a value for  $\alpha_s$ , but the lattice scale must be related to a physical energy/momentum scale. This is achieved with the  $\Upsilon$ - $\Upsilon$  mass difference, however, many other quantities could be used as well [343]. HPQCD obtains  $\alpha_s(M_Z^2) = 0.1184 \pm 0.0006$ , where the uncertainty includes effects from truncating perturbation theory, finite lattice spacing and extrapolation of lattice data. An independent perturbative analysis of a subset of the same lattice-QCD data yields  $\alpha_s(M_Z^2) = 0.1192 \pm 0.0011$  [344]. Using another, independent methodology, the current-current correlator method, HPQCD obtains  $\alpha_s(M_Z^2) = 0.1182 \pm 0.0007$  [342,345]. The analysis of Ref. 346, which uses the Schrödinger functional scheme and avoids the staggered fermion treatment of Ref. 342, finds  $\alpha_s(M_Z^2) = 0.1205 \pm 0.0008 \pm 0.0005^{+0.0000}_{-0.0017}$ , where the first uncertainty is statistical and the others are from systematics. Since this approach uses a different discretization of lattice fermions and a different general methodology, it provides an independent cross check of other lattice extractions of  $\alpha_s$ . A study of the ETM collaboration [347] used lattice data with  $u, d, s$  and  $c$  quarks in the sea and examined the ghost-gluon coupling, obtaining  $\alpha_s(M_Z^2) = 0.1196 \pm 0.0012$ . Finally, a determination of  $\alpha_s$  from the



QCD static potential [348] results in  $\alpha_s(M_Z^2) = 0.1166^{+0.0012}_{-0.0008}$ . The JLQCD collaboration, in an analysis of Adler functions, has recently corrected their initial result of  $\alpha_s(M_Z^2) = 0.1181^{+0.0014}_{-0.0012}$  downwards, by more than 5 standard deviations of their assigned uncertainty, to  $\alpha_s(M_Z^2) = 0.1118^{+0.0016}_{-0.0017}$  [349]. For this and other reasons discussed in [350], we do not include this result in our determination of the average lattice result.

A summary of the results discussed above is given in Fig. 9.2. They average, applying the method of taking the unweighted averages of the central values and their quoted uncertainties at face value, to  $\alpha_s(M_Z^2) = 0.1188 \pm 0.0011$ , which we take as our second result for the determination of the world average value of  $\alpha_s$ . This compares well to a similar compilation and summary provided by the FLAG Working Group [350], suggesting  $\alpha_s(M_Z^2) = 0.1184 \pm 0.0012$  as the overall average of lattice determinations of  $\alpha_s$ . Both these error estimates are more conservative than the one ( $\pm 0.0005$ ) we used in our previous *Review* where we applied the  $\chi^2$  averaging method.

#### 9.4.3. Deep inelastic lepton-nucleon scattering (DIS) :

Studies of DIS final states have led to a number of precise determinations of  $\alpha_s$ : a combination [351] of precision measurements at HERA, based on NLO fits to inclusive jet cross sections in neutral current DIS at high  $Q^2$ , provides combined values of  $\alpha_s$  at different energy scales  $Q$ , as shown in Fig. 9.3, and quotes a combined result of  $\alpha_s(M_Z^2) = 0.1198 \pm 0.0032$ . A more recent study of multijet production [352], based on improved reconstruction and data calibration, confirms the general picture, albeit with a somewhat smaller value of  $\alpha_s(M_Z^2) = 0.1165 \pm 0.0039$ , still in NLO. An evaluation of inclusive jet production, including *approximate* NNLO contributions [353], reduces the theoretical prediction for jet production in DIS, improves the description of the final HERA data in particular at high photon virtuality  $Q^2$  and increases the central fit value of the strong coupling constant.

Another class of studies, analyzing structure functions in NNLO QCD (and partly beyond), provide results which serve as relevant inputs for the world average of  $\alpha_s$ . Most of these studies do *not*, however, explicitly include estimates of theoretical uncertainties when quoting fit results of  $\alpha_s$ . In such cases we add, in quadrature, half of the difference between the results obtained in NNLO and NLO to the quoted errors: A combined analysis of non-singlet structure functions from DIS [354], based on QCD predictions up to N<sup>3</sup>LO in some of its parts, results in  $\alpha_s(M_Z^2) = 0.1141 \pm 0.0022$  (BBG). Studies of singlet and non-singlet structure functions, based on NNLO predictions, result in  $\alpha_s(M_Z^2) = 0.1134 \pm 0.0025$  [355] (ABM) and in  $\alpha_s(M_Z^2) = 0.1158 \pm 0.0036$  [356] (JR). The MSTW group [357], also including data on jet production at the Tevatron, obtains, at NNLO<sup>‡</sup>,  $\alpha_s(M_Z^2) = 0.1171 \pm 0.0024$ . A recent update of this analysis, also including hadron collider data, determined a new set of parton density functions (MMHT2014) [358], together with  $\alpha_s(M_Z^2) = 0.1172 \pm 0.0013$ . The NNPDF group [359] presented a result,  $\alpha_s(M_Z^2) = 0.1173 \pm 0.0011$ , which is in line with the one from the MMHT group, including rather small experimental and theoretical uncertainties of only 6 and 9 per-mille, respectively.

We note that criticism has been expressed on some of the above extractions. Among the issues raised, we mention the neglect of singlet contributions at  $x \geq 0.3$  in pure non-singlet fits [360], the impact and detailed treatment of particular classes of data in the fits [360,361], possible biases due to insufficiently flexible parametrizations of the PDFs [362] and the use of a fixed-flavor number scheme [363,364].

Summarizing the results from world data on structure functions, taking the *unweighted average* of the central values and errors of all selected results, leads to a pre-average value of  $\alpha_s(M_Z^2) = 0.1156 \pm 0.0021$ , see Fig. 9.2.

<sup>‡</sup> Note that for jet production at a hadron collider, only NLO predictions are available, while for the structure functions full NNLO was utilized.

#### 9.4.4. Heavy quarkonia decays :

The most recent extraction of the strong coupling constant from an analysis of radiative  $\Upsilon$  decays [365] resulted in  $\alpha_s(M_Z^2) = 0.119^{+0.006}_{-0.005}$ . This determination is based on QCD at NLO only, so it will not be considered for the final extraction of the world average value of  $\alpha_s$ ; it is, however, an important ingredient for the demonstration of Asymptotic Freedom as given in Fig. 9.3.

#### 9.4.5. Hadronic final states of $e^+e^-$ annihilations :

Re-analyses of event shapes in  $e^+e^-$  annihilation, measured around the  $Z$  peak and at LEP2 center-of-mass energies up to 209 GeV, using NNLO predictions matched to NLL resummation and Monte Carlo models to correct for hadronization effects, resulted in  $\alpha_s(M_Z^2) = 0.1224 \pm 0.0039$  (ALEPH) [366], with a dominant theoretical uncertainty of 0.0035, and in  $\alpha_s(M_Z^2) = 0.1189 \pm 0.0043$  (OPAL) [367]. Similarly, an analysis of JADE data [368] at center-of-mass energies between 14 and 46 GeV gives  $\alpha_s(M_Z^2) = 0.1172 \pm 0.0051$ , with contributions from hadronization model and from perturbative QCD uncertainties of 0.0035 and 0.0030, respectively. Precise determinations of  $\alpha_s$  from 3-jet production alone, in NNLO, resulted in  $\alpha_s(M_Z^2) = 0.1175 \pm 0.0025$  [369] from ALEPH data and in  $\alpha_s(M_Z^2) = 0.1199 \pm 0.0059$  [370] from JADE. These results are summarized in the upper half of Fig. 9.2(d).

Another class of  $\alpha_s$  determinations is based on analytic calculations of non-perturbative and hadronization effects, rather than on Monte Carlo models [371–374], using methods like power corrections, factorization of soft-collinear effective field theory, dispersive models and low scale QCD effective couplings. In these studies, the world data on Thrust distributions, or - most recently - C-parameter distributions, are analysed and fitted to perturbative QCD predictions in NNLO matched with resummation of leading logs up to N<sup>3</sup>LL accuracy, see Sec. 9.2.3.3. The results are  $\alpha_s(M_Z^2) = 0.1164^{+0.0028}_{-0.0024}$  [371],  $\alpha_s(M_Z^2) = 0.1135 \pm 0.0011$  [372] and  $\alpha_s(M_Z^2) = 0.1137^{+0.0034}_{-0.0027}$  [373] from Thrust, and  $\alpha_s(M_Z^2) = 0.1123 \pm 0.0015$  [374] from C-parameter. They are also displayed in Fig. 9.2.

Not to be included in the computation of the world average but worth mentioning are a computation of the NLO corrections to 5-jet production and comparison to the measured 5-jet rates at LEP [375], giving  $\alpha_s(M_Z^2) = 0.1156^{+0.0041}_{-0.0034}$ , and a computation of non-perturbative and perturbative QCD contributions to the scale evolution of quark and gluon jet multiplicities, including resummation, resulting in  $\alpha_s(M_Z^2) = 0.1199 \pm 0.0026$  [376].

We note that there is criticism on both classes of  $\alpha_s$  extractions described above: those based on corrections of non-perturbative hadronization effects using QCD-inspired Monte Carlo generators (since the parton level of a Monte Carlo simulation is not defined in a manner equivalent to that of a fixed-order calculation), as well as studies based on non-perturbative analytic calculations, as their systematics have not yet been fully verified. In particular, quoting rather small overall experimental, hadronization and theoretical uncertainties of only 2, 5 and 9 per-mille, respectively [372,374], seems unrealistic and has neither been met nor supported by other authors or groups.

In view of these open questions, the determination of the *unweighted average* and uncertainties is supposed to provide the most appropriate and unbiased estimate of the average value of  $\alpha_s(M_Z^2)$  for this sub-field, which results in  $\alpha_s(M_Z^2) = 0.1169 \pm 0.0034$ .

#### 9.4.6. Hadron collider results :

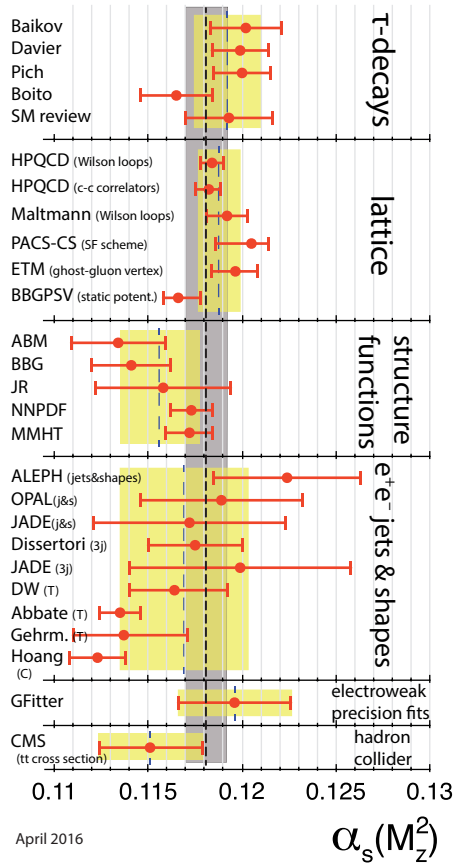
Significant determinations of  $\alpha_s$  from data at hadron colliders, *i.e.* the Tevatron and the LHC, are obtained, however mostly still limited to QCD at NLO. At  $\sqrt{s} = 1.96$  TeV,

$$\alpha_s(M_Z^2) = 0.1161^{+0.0041}_{-0.0048} \text{ and } \alpha_s(M_Z^2) = 0.1191^{+0.0048}_{-0.0071}$$

result from studies of inclusive jet cross sections [377] and from jet angular correlations [378], respectively. ATLAS data on inclusive jet production at  $\sqrt{s} = 7$  TeV [379] lead to [380]

$$\alpha_s(M_Z^2) = 0.1151^{+0.0093}_{-0.0087}.$$





**Figure 9.2:** Summary of determinations of  $\alpha_s(M_Z^2)$  from the six sub-fields discussed in the text. The yellow (light shaded) bands and dashed lines indicate the pre-average values of each sub-field. The dotted line and grey (dark shaded) band represent the final world average value of  $\alpha_s(M_Z^2)$ .

Here, experimental systematics, the choice of jet scale and the use of different PDFs dominate the large overall uncertainties. Determinations of  $\alpha_s$  from CMS data on the ratio of inclusive 3-jet to 2-jet cross sections [381], from inclusive jet production [382] and from the 3-jet differential cross section [383] quoted values of

$$\begin{aligned}\alpha_s(M_Z^2) &= 0.1148 \pm 0.0014(\text{exp.})^{+0.0053}_{-0.0023}(\text{theo.}) , \\ \alpha_s(M_Z^2) &= 0.1185 \pm 0.0019(\text{exp.})^{+0.0060}_{-0.0037}(\text{theo.}) \text{ and} \\ \alpha_s(M_Z^2) &= 0.1171 \pm 0.0013(\text{exp.})^{+0.0073}_{-0.0047}(\text{theo.}) ,\end{aligned}$$

respectively. Most recently, the ATLAS collaboration reported

$$\begin{aligned}\alpha_s(M_Z^2) &= 0.1173 \pm 0.0010(\text{exp.})^{+0.0065}_{-0.0026}(\text{theo.}) \text{ and} \\ \alpha_s(M_Z^2) &= 0.1195 \pm 0.0018(\text{exp.})^{+0.0062}_{-0.0022}(\text{theo.})\end{aligned}$$

using the transverse energy-energy correlation function (TEEC) and its associated azimuthal asymmetry (ATEEC), respectively [247]. All these results are at NLO only, however they provide valuable new values of  $\alpha_s$  at energy scales now extending up to 1.4 TeV. Although not contributing to the overall world average of  $\alpha_s$  which we determine below, it may be worth mentioning that the collider results listed above average to a value of  $\alpha_s(M_Z^2) = 0.1172 \pm 0.0059$ .

So far, only one analysis is available which involves the determination of  $\alpha_s$  from hadron collider data in NNLO of QCD: from a measurement of the  $t\bar{t}$  cross section at  $\sqrt{s} = 7$  TeV, CMS [325] determined

$$\alpha_s(M_Z^2) = 0.1151^{+0.0028}_{-0.0027} ,$$

whereby the dominating contributions to the overall error are experimental ( $^{+0.0017}_{-0.0018}$ ), from parton density functions ( $^{+0.0013}_{-0.0011}$ ) and the value of the top quark pole mass ( $\pm 0.0013$ ).

This latter result will enter our determination of the new world average of  $\alpha_s$ , and will thereby open a new sub-field of  $\alpha_s$  determinations in this *Review*. We note, however, that so far there is only this one result in this sub-field. While there are more recent measurements of  $t\bar{t}$  cross sections from ATLAS and from CMS, at  $\sqrt{s} = 7, 8$  and at 13 TeV, none quotes further extractions of  $\alpha_s$ . A more reliable result will thus be left to the next *Review*, however we note that the most recent measurements of  $t\bar{t}$  cross sections imply larger values of  $\alpha_s(M_Z^2)$  than the one which we use, at this time, as result for this sub-field.

#### 9.4.7. Electroweak precision fit :

The N<sup>3</sup>LO calculation of the hadronic Z decay width [28] was used in the latest update of the global fit to electroweak precision data [384], resulting in

$$\alpha_s(M_Z^2) = 0.1196 \pm 0.0030 ,$$

claiming a negligible theoretical uncertainty. We note that results from electroweak precision data, however, strongly depend on the strict validity of Standard Model predictions and the existence of the minimal Higgs mechanism to implement electroweak symmetry breaking. Any - even small - deviation of nature from this model could strongly influence this extraction of  $\alpha_s$ .

#### 9.4.8. Determination of the world average value of $\alpha_s(M_Z^2)$ :

Obtaining a world average value for  $\alpha_s(M_Z^2)$  is a non-trivial exercise. A certain arbitrariness and subjective component is inevitable because of the choice of measurements to be included in the average, the treatment of (non-Gaussian) systematic uncertainties of mostly theoretical nature, as well as the treatment of correlations among the various inputs, of theoretical as well as experimental origin.

We have chosen to determine pre-averages for sub-fields of measurements which are considered to exhibit a maximum of independence between each other, considering experimental as well as theoretical issues. The six pre-averages are summarized in Fig. 9.2. We recall that these are exclusively obtained from extractions which are based on (at least) full NNLO QCD predictions, and are published in peer-reviewed journals at the time of completing this *Review*. These pre-averages are then combined to the final world average value of  $\alpha_s(M_Z^2)$ , using the  $\chi^2$  averaging method and error treatment as described above. From these, we determine the new world average value of

$$\alpha_s(M_Z^2) = 0.1181 \pm 0.0011 , \quad (9.23)$$

with an uncertainty of 0.9 %.\*\*\* This world average value is in reasonable agreement with that from the 2013 version of this *Review*, which was  $\alpha_s(M_Z^2) = 0.1185 \pm 0.0006$ , however at a somewhat decreased central value and with an overall uncertainty that has almost doubled. These changes are mainly due to the following developments:

- the uncertainty of the combined lattice result, now using the same averaging procedure as applied to the other sub-fields, is more conservative than that used in our previous *Review*, leading to a larger final uncertainty of the new world average, and to a reduced fixing power towards the central average value;
- the relatively low value of  $\alpha_s$  from hadron collider results, which currently consists of only one measurement of the  $t\bar{t}$  cross section at  $\sqrt{s} = 7$  TeV [325] that is likely to be a fluctuation to the low side.

For convenience, we also provide the values for  $\Lambda_{\overline{MS}}$  which correspond to the new world average:

$$\Lambda_{\overline{MS}}^{(6)} = (89 \pm 6) \text{ MeV}, \quad (9.24a)$$

\*\*\* The weighted average, treating all inputs as uncorrelated measurements with Gaussian uncertainties, results in  $\alpha_s(M_Z^2) = 0.11810 \pm 0.00078$  with  $\chi^2/\text{d.o.f.} = 3.7/5$ . Requiring  $\chi^2/\text{d.o.f.}$  to reach unity calls for an overall correlation factor of 0.28, which increases the overall uncertainty to  $\pm 0.00114$ .

$$\Lambda_{\overline{MS}}^{(5)} = (210 \pm 14) \text{ MeV}, \quad (9.24b)$$

$$\Lambda_{\overline{MS}}^{(4)} = (292 \pm 16) \text{ MeV}, \quad (9.24c)$$

$$\Lambda_{\overline{MS}}^{(3)} = (332 \pm 17) \text{ MeV}, \quad (9.24d)$$

for  $n_f = 6, 5, 4$  and 3 quark flavors, which are determined using the 4-loop expression for the running of  $\alpha_s$  according to Eq. (9.5) and 3-loop matching at the charm-, bottom- and top-quark pole masses of 1.3, 4.2 and 173 GeV/ $c^2$ , respectively. Note that for scales below a few GeV, Eq. (9.5) starts to differ significantly from the exact solution of the renormalization group equation Eq. (9.3) and the latter is then to be preferred.

In order to further test and verify the sensitivity of the new average value of  $\alpha_s(M_Z^2)$  to the different pre-averages and fields of  $\alpha_s$  determinations, we give each of the averages obtained when leaving out one of the six input values, as well as the respective, initial value of  $\chi^2$ :

$$\alpha_s(M_Z^2) = 0.1179 \pm 0.0011 \quad (\text{w/o } \tau \text{ results}; \quad \chi_0^2/\text{d.o.f.} = 3.3/4), \quad (9.25a)$$

$$\alpha_s(M_Z^2) = 0.1174 \pm 0.0016 \quad (\text{w/o lattice results}; \quad \chi_0^2/\text{d.o.f.} = 2.9/4), \quad (9.25b)$$

$$\alpha_s(M_Z^2) = 0.1185 \pm 0.0013 \quad (\text{w/o DIS results}; \quad \chi_0^2/\text{d.o.f.} = 2.0/4), \quad (9.25c)$$

$$\alpha_s(M_Z^2) = 0.1182 \pm 0.0010 \quad (\text{w/o } e^+e^- \text{ results}; \quad \chi_0^2/\text{d.o.f.} = 3.5/4), \quad (9.25d)$$

$$\alpha_s(M_Z^2) = 0.1184 \pm 0.0012 \quad (\text{w/o hadron collider}; \quad \chi_0^2/\text{d.o.f.} = 2.4/4) \text{ and} \quad (9.25e)$$

$$\alpha_s(M_Z^2) = 0.1180 \pm 0.0010 \quad (\text{w/o e.w. precision fit}; \quad \chi_0^2/\text{d.o.f.} = 3.4/4). \quad (9.25f)$$

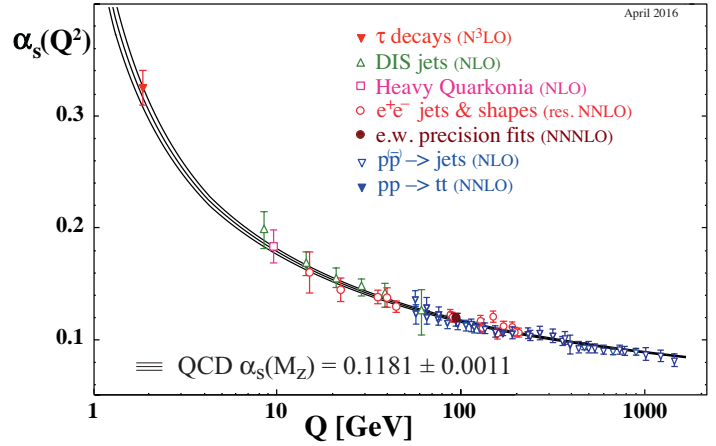
They are well within the uncertainty of the overall world average quoted above. Note, however, that the average *excluding* the lattice result is no longer as close to the value obtained from lattice alone as was the case in the 2013 *Review*, but is now smaller by almost one standard deviation of its assigned uncertainty.

Notwithstanding the many open issues still present within each of the sub-fields summarised in this *Review*, the wealth of available results provides a rather precise and reasonably stable world average value of  $\alpha_s(M_Z^2)$ , as well as a clear signature and proof of the energy dependence of  $\alpha_s$ , in full agreement with the QCD prediction of Asymptotic Freedom. This is demonstrated in Fig. 9.3, where results of  $\alpha_s(Q^2)$  obtained at discrete energy scales  $Q$ , now also including those based just on NLO QCD, are summarized. Thanks to the results from the Tevatron and from the LHC, the energy scales at which  $\alpha_s$  is determined now extend up to more than 1 TeV $^\diamond$ .

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$^\diamond$  We note, however, that in many such studies, like those based on exclusive states of jet multiplicities, the relevant energy scale of the measurement is not uniquely defined. For instance, in studies of the ratio of 3- to 2-jet cross sections at the LHC, the relevant scale was taken to be the average of the transverse momenta of the two leading jets [381], but could alternatively have been chosen to be the transverse momentum of the 3 $^{rd}$  jet.



**Figure 9.3:** Summary of measurements of  $\alpha_s$  as a function of the energy scale  $Q$ . The respective degree of QCD perturbation theory used in the extraction of  $\alpha_s$  is indicated in brackets (NLO: next-to-leading order; NNLO: next-to-next-to leading order; res. NNLO: NNLO matched with resummed next-to-leading logs; NNNLO (N<sup>3</sup>LO): next-to-NNLO).

## References:

1. R.K. Ellis, W.J. Stirling, and B.R. Webber, “*QCD and collider physics*,” Camb. Monogr. Part. Phys. Nucl. Phys. Cosmol. **81** (1996).
2. C.A. Baker *et al.*, Phys. Rev. Lett. **97**, 131801 (2006).
3. H.-Y. Cheng, Phys. Reports **158**, 1 (1988).
4. G. Dissertori, I.G. Knowles, and M. Schmelling, “*High energy experiments and theory*,” Oxford, UK: Clarendon (2003).
5. R. Brock *et al.*, [CTEQ Collab.], Rev. Mod. Phys. **67**, 157 (1995), see also <http://www.phys.psu.edu/~cteq/handbook/v1.1/handbook.pdf>.
6. A.S. Kronfeld and C. Quigg, Am. J. Phys. **78**, 1081 (2010).
7. T. Plehn, Lect. Notes Phys. **844**, 1 (2012).
8. R. Stock (Ed.), Relativistic Heavy Ion Physics, Springer-Verlag Berlin, Heidelberg, 2010.
9. *Proceedings of the XXIV International Conference on Ultrarelativistic Nucleus–Nucleus Collisions, Quark Matter 2014*, Nucl. Phys. A, volume 931.
10. R. Hwa and X. N. Wang (Ed.), Quark Gluon Plasma 4, World Scientific Publishing Company, 2010.
11. T. van Ritbergen, J.A.M. Vermaseren, and S.A. Larin, Phys. Lett. **B400**, 379 (1997).
12. M. Czakon, Nucl. Phys. **B710**, 485 (2005).
13. W.A. Bardeen *et al.*, Phys. Rev. **D18**, 3998 (1978).
14. D.J. Gross and F. Wilczek, Phys. Rev. Lett. **30**, 1343 (1973).
15. H.D. Politzer, Phys. Rev. Lett. **30**, 1346 (1973).
16. Y. Schroder and M. Steinhauser, JHEP **0601**, 051 (2006).
17. K.G. Chetyrkin, J.H. Kuhn, and C. Sturm, Nucl. Phys. **B744**, 121 (2006).
18. A.G. Grozin *et al.*, JHEP **1109**, 066 (2011).
19. K.G. Chetyrkin, B.A. Kniehl, and M. Steinhauser, Nucl. Phys. **B510**, 61 (1998).
20. K.G. Chetyrkin, J.H. Kuhn, and M. Steinhauser, Comp. Phys. Comm. **133**, 43 (2000).
21. B. Schmidt and M. Steinhauser, Comp. Phys. Comm. **183**, 1845 (2012) <http://www.ttp.kit.edu/Progdata/ttp12/ttp12-02/>.
22. A.V. Bednyakov, Phys. Lett. **B741**, 262 (2015).
23. M. Beneke, Phys. Reports **317**, 1 (1999).
24. P.A. Baikov *et al.*, Phys. Lett. **B714**, 62 (2012).
25. K.G. Chetyrkin, J.H. Kuhn, and A. Kwiatkowski, Phys. Reports **277**, 189 (1996).
26. Y. Kiyo *et al.*, Nucl. Phys. **B823**, 269 (2009).
27. P.A. Baikov *et al.*, Phys. Rev. Lett. **108**, 222003 (2012).

28. P.A. Baikov, K.G. Chetyrkin, and J.H. Kuhn, Phys. Rev. Lett. **101**, 012002 (2008).
29. P.A. Baikov, K.G. Chetyrkin, and J.H. Kuhn, Phys. Rev. Lett. **96**, 012003 (2006).
30. P.A. Baikov and K.G. Chetyrkin, Phys. Rev. Lett. **97**, 061803 (2006).
31. D. Asner *et al.*, arXiv:1307.8265 [hep-ex].
32. H.W. Lin *et al.*, Phys. Rev. **D91**, 054510 (2015);  
C. Alexandrou *et al.*, Phys. Rev. **D92**, 014502 (2015).
33. S. Forte and G. Watt, Ann. Rev. Nucl. and Part. Sci. **63**, 291 (2013).
34. J. D. Bjorken and E. A. Paschos, Phys. Rev. **185**, 1975 (1969).
35. J.A.M. Vermaseren, A. Vogt, and S. Moch, Nucl. Phys. **B724**, 3 (2005).
36. E.B. Zijlstra and W.L. van Neerven, Phys. Lett. **B297**, 377 (1992).
37. S. Moch, J.A.M. Vermaseren, and A. Vogt, Nucl. Phys. **B813**, 220 (2009).
38. E. Laenen *et al.*, Nucl. Phys. **B392**, 162 (1993);  
S. Riemersma, J. Smith, and W.L. van Neerven, Phys. Lett. **B347**, 143 (1995).
39. J. Blümlein, A. De Freitas, and C. Schneider, Nucl. Part. Phys. Proc. **261-262** 185 Nucl. Part. Phys. Proc. **261-262** (2015) 185.
40. J.C. Collins, D.E. Soper, and G.F. Sterman, Adv. Ser. Direct. High Energy Phys. **5**, 1 (1988).
41. J.C. Collins, *Foundations of Perturbative QCD*, Cambridge University Press, 2011.
42. G.C. Nayak, J.-W. Qiu, and G.F. Sterman, Phys. Rev. **D72**, 114012 (2005).
43. V.N. Gribov and L.N. Lipatov, Sov. J. Nucl. Phys. **15**, 438 (1972);  
L.N. Lipatov, Sov. J. Nucl. Phys. **20**, 94 (1975);  
G. Altarelli and G. Parisi, Nucl. Phys. **B126**, 298 (1977);  
Yu.L. Dokshitzer, Sov. Phys. JETP **46**, 641 (1977).
44. G. Curci, W. Furmanski, and R. Petronzio, Nucl. Phys. **B175**, 27 (1980);  
W. Furmanski and R. Petronzio, Phys. Lett. **B97**, 437 (1980).
45. A. Vogt, S. Moch, and J.A.M. Vermaseren, Nucl. Phys. **B691**, 129 (2004);  
S. Moch, J.A.M. Vermaseren, and A. Vogt, Nucl. Phys. **B688**, 101 (2004).
46. R.S. Thorne, Phys. Rev. **D73**, 054019 (2006).
47. S. Forte *et al.*, Nucl. Phys. **B834**, 116 (2010).
48. M. Guzzi *et al.*, Phys. Rev. **D86**, 053005 (2012).
49. J.C. Collins, D.E. Soper, and G. Sterman, Nucl. Phys. **B261**, 104 (1985).
50. R. Hamberg, W.L. van Neerven, and T. Matsuura, Nucl. Phys. **B359**, 343 (1991); Erratum *ibid.*, **B 644** 403, (2002).
51. R.V. Harlander and W.B. Kilgore, Phys. Rev. Lett. **88**, 201801 (2002).
52. O. Brein, A. Djouadi, and R. Harlander, Phys. Lett. **B579**, 149 (2004).
53. P. Bolzoni *et al.*, Phys. Rev. Lett. **105**, 011801 (2010).
54. D. de Florian and J. Mazzitelli, Phys. Rev. Lett. **111**, 201801 (2013).
55. M. Czakon, P. Fiedler, and A. Mitov, Phys. Rev. Lett. **110**, 252004 (2013).
56. T. Gehrmann *et al.*, Phys. Rev. Lett. **113**, 21 (2014).
57. F. Cascioli *et al.*, Phys. Lett. **B735**, 311 (2014).
58. C. Anastasiou *et al.*, Phys. Rev. Lett. **114**, 21 (2015).
59. S. Dittmaier *et al.*, [LHC Higgs Cross Section Working Group], arXiv:1101.0593 [hep-ph].
60. M. Greco and A. Vicini, Nucl. Phys. **B415**, 386 (1994).
61. L.N. Lipatov, Sov. J. Nucl. Phys. **23**, 338 (1976) [Yad. Fiz. **23**, 642 (1976)].
62. E.A. Kuraev, L.N. Lipatov, and V.S. Fadin, Sov. Phys. JETP **45**, 199 (1977) [Zh. Eksp. Teor. Fiz. **72**, 377 (1977)].
63. I.I. Balitsky and L.N. Lipatov, Sov. J. Nucl. Phys. **28**, 822 (1978) [Yad. Fiz. **28**, 1597 (1978)].
64. V.S. Fadin and L.N. Lipatov, Phys. Lett. **B429**, 127 (1998).
65. M. Ciafaloni and G. Camici, Phys. Lett. **B430**, 329 (1998).
66. G. Altarelli, R.D. Ball, and S. Forte, Nucl. Phys. **B799**, 199 (2008).
67. M. Ciafaloni *et al.*, JHEP **0708**, 046 (2007).
68. C.D. White and R.S. Thorne, Phys. Rev. **D75**, 034005 (2007).
69. E. Iancu *et al.*, Phys. Lett. **B744**, 293 (2015).
70. N. Gromov, F. Levkovich-Maslyuk, and G. Sizov, arXiv:1507.04010 [hep-th].
71. V. N. Velizhanin, arXiv:1508.02857 [hep-th].
72. I. Balitsky, Nucl. Phys. **B463**, 99 (1996).
73. Y.V. Kovchegov, Phys. Rev. **D60**, 034008 (1999).
74. A. Hebecker, Phys. Reports **331**, 1 (2000).
75. A.V. Belitsky and A.V. Radyushkin, Phys. Reports **418**, 1 (2005).
76. E. Boos *et al.*, [CompHEP Collab.], Nucl. Instrum. Methods **A534**, 250 (2004) <http://comphep.sinp.msu.ru/>.
77. J. Alwall *et al.*, JHEP **1407**, 079 (2014), <https://launchpad.net/mg5amcno>.
78. M.L. Mangano *et al.*, JHEP **0307**, 001 (2003), <http://cern.ch/mlm/alpgen/>.
79. T. Gleisberg and S. Hoche, JHEP **0812**, 039 (2008), <https://sherpa.hepforge.org/trac/wiki>.
80. A. Cafarella, C.G. Papadopoulos, and M. Worek, Comp. Phys. Comm. **180**, 1941 (2009), <http://cern.ch/helac-phegas/>.
81. F.A. Berends and W.T. Giele, Nucl. Phys. **B306**, 759 (1988).
82. L.J. Dixon, arXiv:hep-ph/9601359.
83. Z. Bern, L.J. Dixon, and D.A. Kosower, Ann. Phys. **322**, 1587 (2007).
84. S. Badger *et al.*, Phys. Rev. **D87**, 3 (2013).
85. S. Catani and M.H. Seymour, Nucl. Phys. **B485**, 291 (1997) [Erratum-*ibid.* **B 510**, 503 (1998)].
86. S. Frixione, Z. Kunszt, and A. Signer, Nucl. Phys. **B467**, 399 (1996).
87. D.A. Kosower, Phys. Rev. **D57**, 5410 (1998);  
J.M. Campbell, M.A. Cullen, and E.W.N. Glover, Eur. Phys. J. **C9**, 245 (1999);  
D.A. Kosower, Phys. Rev. **D71**, 045016 (2005).
88. Z. Nagy, Phys. Rev. **D68**, 094002 (2003), <http://www.desy.de/~znagy/Website/NLOJet++.html>.
89. J.M. Campbell and R.K. Ellis, Phys. Rev. **D62**, 114012 (2000).
90. K. Arnold *et al.*, arXiv:1107.4038 [hep-ph]; <http://www-itp.particle.uni-karlsruhe.de/~vbfnlweb/>.
91. T. Binoth *et al.*, Eur. Phys. J. **C16**, 311 (2000), [http://lapth.in2p3.fr/PHOX\\_FAMILY/](http://lapth.in2p3.fr/PHOX_FAMILY/).
92. G. Bevilacqua *et al.*, Comp. Phys. Comm. **184**, 986 (2013), <http://cern.ch/helac-phegas/>.
93. G. Cullen *et al.*, Eur. Phys. J. **C74**, 8 (2014), <http://gosam.hepforge.org/>.
94. S. Badger *et al.*, Comp. Phys. Comm. **184**, 1981 (2013), <https://bitbucket.org/njet/njet/wiki/Home>.
95. F. Cascioli, P. Maierhofer, and S. Pozzorini, Phys. Rev. Lett. **108**, 111601 (2012), <https://openloops.hepforge.org/>.
96. T. Gleisberg *et al.*, JHEP **0902**, 007 (2009), <http://projects.hepforge.org/sherpa/>.
97. Z. Bern *et al.*, PoS LL **2012**, 018 (2012).
98. S. Actis *et al.*, JHEP **1304**, 037 (2013).
99. Z. Bern *et al.*, Comp. Phys. Comm. **185**, 1443 (2014).
100. T. Kluge, K. Rabbertz, and M. Wobisch, arXiv:hep-ph/0609285, <http://fastnlo.hepforge.org/>.
101. T. Carli *et al.*, Eur. Phys. J. **C66**, 503 (2010), <https://applgrid.hepforge.org/>.
102. L. Del Debbio, N.P. Hartland, and S. Schumann, Comp. Phys. Comm. **185**, 2115 (2014), <http://mcgrid.hepforge.org/>.
103. V. Bertone *et al.*, JHEP **1408**, 166 (2014), <https://amcfast.hepforge.org/>.
104. G. Cullen, N. Greiner, and G. Heinrich, Eur. Phys. J. **C73**, 4 (2013).
105. S. Kallweit *et al.*, JHEP **1504**, 012 (2015).
106. A. Denner *et al.*, JHEP **1504**, 018 (2015).

107. S. Frixione *et al.* JHEP **1506**, 184 (2015).
108. R.K. Ellis *et al.*, Phys. Reports **518**, 141 (2012).
109. S. Becker *et al.*, Phys. Rev. Lett. **108**, 032005 (2012).
110. Z. Bern *et al.*, Phys. Rev. **D88**, 014025 (2013).
111. S. Badger *et al.*, Phys. Rev. **D89**, 3 (2014).
112. Z. Bern *et al.*, Nucl. Phys. **B425**, 217 (1994).
113. J.M. Campbell and E.W.N. Glover, Nucl. Phys. **B527**, 264 (1998).
114. S. Catani and M. Grazzini, Phys. Lett. **B446**, 143 (1999).
115. T. Binoth and G. Heinrich, Nucl. Phys. **B585**, 741 (2000).
116. C. Anastasiou, K. Melnikov, and F. Petriello, Phys. Rev. **D69**, 076010 (2004).
117. A. Gehrmann-De Ridder, T. Gehrmann, and E.W.N. Glover, JHEP **0509**, 056 (2005).
118. G. Somogyi, Z. Trocsanyi, and V. Del Duca, JHEP **0701**, 070 (2007).
119. M. Czakon, Phys. Lett. **B693**, 259 (2010).
120. S. Catani and M. Grazzini, Phys. Rev. Lett. **98**, 222002 (2007), <http://theory.fi.infn.it/grazzini/codes.html>.
121. R. Boughezal *et al.* Phys. Rev. Lett. **115**, 6 (2015).
122. J. Gaunt *et al.* arXiv:1505.04794 [hep-ph].
123. M. Cacciari *et al.* arXiv:1506.02660 [hep-ph].
124. A. Gehrmann-De Ridder *et al.*, Phys. Rev. Lett. **99**, 132002 (2007); JHEP **0712**, 094 (2007); Phys. Rev. Lett. **100**, 172001 (2008).
125. A. Gehrmann-De Ridder *et al.* Comp. Phys. Comm. **185**, 3331 (2014), <https://eera3.hepforge.org/>.
126. S. Weinzierl, Phys. Rev. Lett. **101**, 162001 (2008); JHEP **0906**, 041 (2009).
127. K. Melnikov and F. Petriello, Phys. Rev. **D74**, 114017 (2006), <http://gate.hep.anl.gov/fpetriello/FEWZ.html>.
128. S. Catani *et al.*, Phys. Rev. Lett. **103**, 082001 (2009), <http://theory.fi.infn.it/grazzini/dy.html>.
129. C. Anastasiou, K. Melnikov, and F. Petriello, Nucl. Phys. **B724**, 197 (2005), <http://www.phys.ethz.ch/~pheno/fehipro/>.
130. G. Ferrera, M. Grazzini, and F. Tramontano, Phys. Rev. Lett. **107**, 152003 (2011).
131. G. Ferrera, M. Grazzini, and F. Tramontano, Phys. Lett. **B740**, 51 (2015).
132. S. Catani *et al.*, Phys. Rev. Lett. **108**, 072001 (2012).
133. M. Grazzini *et al.*, Phys. Lett. **B731**, 204 (2014).
134. M. Grazzini, S. Kallweit, and D. Rathlev, JHEP **1507**, 085 (2015).
135. R. Boughezal *et al.*, arXiv:1504.07922 [hep-ph].
136. R. Boughezal *et al.*, Phys. Lett. **B748**, 5 (2015).
137. A. Gehrmann-De Ridder *et al.*, arXiv:1507.02850 [hep-ph].
138. M. Brucherseifer, F. Caola, and K. Melnikov, Phys. Lett. **B736**, 58 (2014).
139. M. Czakon, P. Fiedler, and A. Mitov, Phys. Rev. Lett. **115**, 5 (2015).
140. J. Currie *et al.*, JHEP **1401**, 110 (2014).
141. Y.L. Dokshitzer, D. Diakonov, and S.I. Troian, Phys. Reports **58**, 269 (1980).
142. G. Parisi and R. Petronzio, Nucl. Phys. **B154**, 427 (1979).
143. G. Curci, M. Greco, and Y. Srivastava, Nucl. Phys. **B159**, 451 (1979).
144. A. Bassetto, M. Ciafaloni, and G. Marchesini, Nucl. Phys. **B163**, 477 (1980).
145. J.C. Collins and D.E. Soper, Nucl. Phys. **B193**, 381 (1981) [Erratum-*ibid.* **B213**, 545 (1983)].
146. J.C. Collins and D.E. Soper, Nucl. Phys. **B197**, 446 (1982).
147. J. Kodaira and L. Trentadue, Phys. Lett. **B112**, 66 (1982).
148. J. Kodaira and L. Trentadue, Phys. Lett. **B123**, 335 (1983).
149. J.C. Collins, D.E. Soper, and G. Sterman, Nucl. Phys. **B250**, 199 (1985).
150. S. Catani *et al.*, Nucl. Phys. **B407**, 3 (1993).
151. C. W. Bauer *et al.*, Phys. Rev. **D63**, 114020 (2001).
152. C.W. Bauer, D. Pirjol, and I.W. Stewart, Phys. Rev. **D65**, 054022 (2002).
153. T. Becher, A. Broggio, and A. Ferroglia, arXiv:1410.1892 [hep-ph].
154. S. Catani *et al.*, Phys. Lett. **B269**, 432 (1991).
155. N. Brown and W.J. Stirling, Phys. Lett. **B252**, 657 (1990).
156. W. Bartel *et al.*, [JADE Collab.], Z. Phys. **C33**, 23 (1986).
157. N. Kidonakis, G. Oderda, and G. Sterman, Nucl. Phys. **B531**, 365 (1998).
158. R. Bonciani *et al.*, Phys. Lett. **B575**, 268 (2003).
159. A. Banfi, G.P. Salam, and G. Zanderighi, JHEP **0503**, 073 (2005).
160. D. de Florian and M. Grazzini, Phys. Rev. Lett. **85**, 4678 (2000).
161. G. Bozzi *et al.*, Nucl. Phys. **B737**, 73 (2006), <http://theory.fi.infn.it/grazzini/codes.html>.
162. G. Bozzi *et al.*, Phys. Lett. **B696**, 207 (2011).
163. T. Becher and M. Neubert, Eur. Phys. J. **C71**, 1665 (2011).
164. T. Becher, M. Neubert, and D. Wilhelm, <http://cute.hepforge.org/>.
165. D. de Florian *et al.*, JHEP **1206**, 132 (2012), <http://theory.fi.infn.it/grazzini/codes.html>.
166. C. Balazs and C.P. Yuan, Phys. Rev. **D56**, 5558 (1997).
167. S. Catani *et al.*, arXiv:1507.06937 [hep-ph].
168. A. Banfi *et al.*, Phys. Lett. **B715**, 152 (2012).
169. M. Grazzini *et al.*, arXiv:1507.02565 [hep-ph].
170. D. de Florian and M. Grazzini, Nucl. Phys. **B704**, 387 (2005).
171. T. Becher and G. Bell, JHEP **1211**, 126 (2012).
172. A. Banfi *et al.*, Phys. Rev. Lett. **109**, 202001 (2012); T. Becher, M. Neubert, and L. Rothen, JHEP **1310**, 125 (2013); I.W. Stewart *et al.*, Phys. Rev. **D89**, 5 (2014).
173. I.W. Stewart, F.J. Tackmann, and W.J. Waalewijn, Phys. Rev. Lett. **106**, 032001 (2011).
174. Y.-T. Chien *et al.*, Phys. Rev. **D87**, 014010 (2013); T.T. Jouttenus *et al.*, Phys. Rev. **D88**, 054031 (2013).
175. M. Dasgupta *et al.*, JHEP **1210**, 126 (2012).
176. V. Ahrens *et al.*, JHEP **1009**, 097 (2010).
177. M. Aliev *et al.*, Comp. Phys. Comm. **182**, 1034 (2011).
178. N. Kidonakis, Phys. Rev. **D82**, 114030 (2010).
179. T. Becher, C. Lorentzen, M. D. Schwartz, Phys. Rev. Lett. **108**, 012001 (2012).
180. T. Becher *et al.*, Eur. Phys. J. **C75**, 4 (2015).
181. E. Gerwick *et al.*, JHEP **1502**, 106 (2015).
182. A. Banfi *et al.*, JHEP **1505**, 102 (2015).
183. T. Becher and M.D. Schwartz, JHEP **0807**, 034 (2008).
184. A. H. Hoang *et al.*, Phys. Rev. **D91**, 9 (2015).
185. Y.-T. Chien and M.D. Schwartz, JHEP **1008**, 058 (2010).
186. P.F. Monni, T. Gehrmann, and G. Luisoni, JHEP **1108**, 010 (2011).
187. S. Catani *et al.*, Nucl. Phys. **B888**, 75 (2014).
188. S. Fleming *et al.*, Phys. Rev. **D77**, 074010 (2008).
189. I. Feige *et al.*, Phys. Rev. Lett. **109**, 092001 (2012).
190. M. Dasgupta *et al.*, JHEP **1309**, 029 (2013).
191. A.J. Larkoski *et al.*, JHEP **1405**, 146 (2014).
192. M. Dasgupta, A. Powling, and A. Siodmok, arXiv:1503.01088 [hep-ph].
193. A.J. Larkoski, I. Moulton, and D. Neill, arXiv:1507.03018 [hep-ph].
194. Yu.L. Dokshitzer *et al.*, “Basics of perturbative QCD,” Gif-sur-Yvette, France: Éditions frontières (1991), see also <http://www.lpthe.jussieu.fr/~yuri/BPQCD/cover.html>.
195. T. Sjostrand *et al.*, Comp. Phys. Comm. **135**, 238 (2001).
196. T. Sjostrand, S. Mrenna, and P. Skands, JHEP **0605**, 026 (2006), <http://projects.hepforge.org/pythia6/>.
197. T. Sjöstrand *et al.*, Comp. Phys. Comm. **191**, 159 (2015), <http://home.thep.lu.se/~torbjorn/Pythia.html>.
198. B.R. Webber, Nucl. Phys. **B238**, 492 (1984).

199. G. Corcella *et al.*, JHEP **0101**, 010 (2001), <http://www.hep.phy.cam.ac.uk/theory/webber/Herwig/>.
200. M. Bahr *et al.*, Eur. Phys. J. **C58**, 639 (2008), <http://projects.hepforge.org/herwig/>.
201. L. Lonnblad, Comp. Phys. Comm. **71**, 15, (1992).
202. A. Buckley *et al.*, Phys. Rept. **504**, 145-233 (2011).
203. B. Andersson *et al.*, Phys. Reports **97**, 31 (1983).
204. T. Sjostrand, Nucl. Phys. **B248**, 469 (1984).
205. T. Sjostrand and M. van Zijl, Phys. Rev. **D36**, 2019 (1987).
206. S. Catani *et al.*, JHEP **0111**, 063 (2001).
207. J. Alwall *et al.*, Eur. Phys. J. **C53**, 473 (2008).
208. S. Frixione and B.R. Webber, JHEP **0206**, 029 (2002).
209. P. Nason, JHEP **0411**, 040 (2004).
210. S. Alioli *et al.*, JHEP **1006**, 043 (2010), <http://powhegbox.mib.infn.it/>.
211. S. Hoeche *et al.*, JHEP **1209**, 049 (2012).
212. S. Hoeche *et al.*, JHEP **1304**, 027 (2013);  
R. Frederix and S. Frixione, JHEP **1212**, 061 (2012);  
S. Plätzer, JHEP **1308**, 114 (2013);  
L. Lonnblad and S. Prestel, JHEP **1303**, 166 (2013);  
K. Hamilton *et al.*, JHEP **1305**, 082 (2013).
213. K. Hamilton *et al.*, JHEP **1310**, 222 (2013);  
A. Karlberg, E. Re, and G. Zanderighi, JHEP **1409**, 134 (2014);  
S. Hoeche, Y. Li, and S. Prestel, [arXiv:1507.05325 \[hep-ph\]](https://arxiv.org/abs/1507.05325);  
S. Alioli *et al.*, [arXiv:1508.01475 \[hep-ph\]](https://arxiv.org/abs/1508.01475).
214. P.M. Stevenson, Phys. Lett. **B100**, 61 (1981).
215. P.M. Stevenson, Phys. Rev. **D23**, 2916 (1981).
216. G. Grunberg, Phys. Rev. **D29**, 2315 (1984).
217. S.J. Brodsky, G.P. Lepage, and P.B. Mackenzie, Phys. Rev. **D28**, 228 (1983).
218. M. Cacciari and N. Houdeau, JHEP **1109**, 039 (2011).
219. A. David and G. Passarino, Phys. Lett. **B726**, 266 (2013).
220. E. Bagnaschi *et al.*, JHEP **1502**, 133 (2015).
221. M. Cacciari *et al.*, JHEP **0404**, 068 (2004).
222. M. Dasgupta and G.P. Salam, J. Phys. **G30**, R143 (2004).
223. S. Moretti, L. Lonnblad, and T. Sjostrand, JHEP **9808**, 001 (1998).
224. G.P. Salam, Eur. Phys. J. **C67**, 637 (2010).
225. S.D. Ellis *et al.*, Prog. in Part. Nucl. Phys. **60**, 484 (2008).
226. M. Cacciari, [arXiv:1509.02272 \[hep-ph\]](https://arxiv.org/abs/1509.02272).
227. G.P. Salam and G. Soyez, JHEP **0705**, 086 (2007).
228. S. Catani *et al.*, Nucl. Phys. **B406**, 187 (1993).
229. S.D. Ellis and D.E. Soper, Phys. Rev. **D48**, 3160 (1993).
230. Y.L. Dokshitzer *et al.*, JHEP **9708**, 001 (1997).
231. M. Wobisch and T. Wengler.
232. M. Cacciari, G.P. Salam, and G. Soyez, JHEP **0804**, 063 (2008).
233. S. Bethke *et al.*, Nucl. Phys. **B370**, 310 (1992), Nucl. Phys. **B523**, 681 (1998).
234. M. Cacciari and G.P. Salam, Phys. Lett. **B641**, 57 (2006);  
M. Cacciari, G.P. Salam, and G. Soyez, Eur. Phys. J. **C72**, 1896 (2012) <http://fastjet.fr/>.
235. S. Brandt *et al.*, Phys. Lett. **12**, 57 (1964).
236. E. Farhi, Phys. Rev. Lett. **39**, 1587 (1977).
237. O. Biebel, Phys. Reports **340**, 165 (2001).
238. S. Kluth, Rept. on Prog. in Phys. **69**, 1771 (2006).
239. C.L. Basham *et al.*, Phys. Rev. Lett. **41**, 1585 (1978).
240. A. Ali, E. Pietarinen, and W. J. Stirling, Phys. Lett. **B141**, 447 (1984).
241. I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, Phys. Rev. Lett. **105**, 092002 (2010).
242. A. Banfi, G.P. Salam, and G. Zanderighi, JHEP **0408**, 062 (2004).
243. A. Banfi, G.P. Salam, and G. Zanderighi, JHEP **1006**, 038 (2010).
244. T. Aaltonen *et al.*, [CDF Collab.], Phys. Rev. **D83**, 112007 (2011).
245. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C72**, 2211 (2012).
246. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D88**, 032004 (2013).
247. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B750**, 427 (2015).
248. V. Khachatryan *et al.*, [CMS Collab.], Phys. Lett. **B699**, 48 (2011).
249. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B722**, 238 (2013).
250. V. Khachatryan *et al.*, [CMS Collab.], JHEP **1410**, 87 (2014).
251. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B730**, 243 (2014).
252. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D83**, 052003 (2011).
253. S. Chatrchyan *et al.*, [CMS Collab.], JHEP **1206**, 160 (2012).
254. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C73**, 12 (2013).
255. C. Glasman [H1 Collab. and ZEUS Collab.], Nucl. Phys. (Proc. Supp.) **191**, 121 (2009).
256. T. Carli, K. Rabbertz, and S. Schumann, [arXiv:1506.03239 \[hep-ex\]](https://arxiv.org/abs/1506.03239).
257. A. Abdesselam *et al.*, Eur. Phys. J. **C71**, 1661 (2011).
258. A. Altheimer *et al.*, J. Phys. **G39**, 063001 (2012).
259. A. Altheimer *et al.*, Eur. Phys. J. **C74**, 3 (2014).
260. S. Schaetzel, Eur. Phys. J. **C75**, 9 (2015).
261. T. Schorner-Sadenius, Eur. Phys. J. **C72**, 2060 (2012).
262. J.M. Campbell, J.W. Huston, and W.J. Stirling, Rept. on Prog. in Phys. **70**, 89 (2007).
263. M.L. Mangano, Phys. Usp. **53**, 109 (2010).
264. J.M. Butterworth, G. Dissertori, and G.P. Salam, Ann. Rev. Nucl. and Part. Sci. **62**, 387 (2012).
265. T. Carli, T. Gehrmann, and S. Hoeche, Eur. Phys. J. **C67**, 73 (2010).
266. F.D. Aaron *et al.*, [H1 Collab.], Eur. Phys. J. **C65**, 363 (2010).
267. F.D. Aaron *et al.*, [H1 Collab.], Eur. Phys. J. **C54**, 389 (2008).
268. S. Chekanov *et al.*, [ZEUS Collab.], Eur. Phys. J. **C52**, 515 (2007).
269. S. Chekanov *et al.*, [ZEUS Collab.], Phys. Rev. **D78**, 032004 (2008).
270. H. Abramowicz *et al.*, [ZEUS Collab.], Eur. Phys. J. **C70**, 965 (2010).
271. H. Abramowicz *et al.*, [ZEUS Collab.], Phys. Lett. **B691**, 127 (2010).
272. S. Chekanov *et al.*, [ZEUS Collab.], Phys. Rev. **D85**, 052008 (2012).
273. F.D. Aaron *et al.*, [H1 Collab.], Eur. Phys. J. **C67**, 1 (2010).
274. V. Andreev *et al.*, [H1 Collab.], Eur. Phys. J. **C75**, 2 (2015).
275. S. Chekanov *et al.*, [ZEUS Collab.], Nucl. Phys. **B792**, 1 (2008).
276. S. Chekanov *et al.*, [ZEUS Collab.], Phys. Rev. **D76**, 072011 (2007).
277. A. Aktas *et al.*, [H1 Collab.], Phys. Lett. **B639**, 21 (2006).
278. H. Abramowicz *et al.*, [ZEUS Collab.], Eur. Phys. J. **C71**, 1659 (2011).
279. H. Abramowicz *et al.*, [ZEUS Collab.], Nucl. Phys. **B864**, 1 (2012).
280. <http://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/CombinedSummaryPlots/SM>.
281. <http://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsCombined>.
282. A. Abulencia *et al.*, [CDF - Run II Collab.], Phys. Rev. **D75**, 092006 (2007) [Erratum-ibid. 119901].
283. V.M. Abazov *et al.*, [D0 Collab.], Phys. Rev. Lett. **101**, 062001 (2008).
284. V.M. Abazov *et al.*, [D0 Collab.], Phys. Rev. **D85**, 052006 (2012).
285. B. Abelev *et al.*, [ALICE Collab.], Phys. Lett. **B722**, 262 (2013).
286. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C73**, 2509 (2013).
287. G. Aad *et al.*, [ATLAS Collab.], JHEP **1502**, 153 (2015).
288. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Rev. **D87**, 112002 (2013).
289. M. Wobisch *et al.*, [fastNLO Collab.], [arXiv:1109.1310 \[hep-ph\]](https://arxiv.org/abs/1109.1310).
290. P. Francavilla, article submitted to the International Journal of Modern Physics A (IJMPA) as part of the special issue on the "Jet Measurements at the LHC", editor G. Dissertori, 2015.

291. “Plots of Cross Sections and Related Quantities” review in this edition of *Review* [pdg.lbl.gov].
292. A. Schwartzman, [arXiv:1509.05459 [hep-ex]].
293. J. Rojo, *Int. J. Mod. Phys. A* **30**, 1546005 (2015).
294. G. Aad *et al.*, [ATLAS Collab.], *JHEP* **1405**, 059 (2014).
295. T. Aaltonen *et al.*, [CDF Collab.], *Phys. Rev. D* **79**, 112002 (2009).
296. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Rev. Lett.* **103**, 191803 (2009).
297. S. Chatrchyan *et al.*, [CMS Collab.], *JHEP* **1205**, 055 (2012).
298. V. Khachatryan *et al.*, [CMS Collab.], *Phys. Lett. B* **746**, 79 (2015).
299. G. Aad *et al.*, [ATLAS Collab.], *JHEP* **1301**, 029 (2013).
300. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Rev. Lett.* **94**, 221801 (2005).
301. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Lett. B* **721**, 212 (2013).
302. G. Aad *et al.*, [ATLAS Collab.], *Phys. Rev. Lett.* **106**, 172002 (2011).
303. V. Khachatryan *et al.*, [CMS Collab.], *Phys. Rev. Lett.* **106**, 122003 (2011).
304. M. Campanelli, arXiv:1508.04631 [hep-ph].
305. P. Kokkas, arXiv:1509.02144 [hep-ex].
306. G. Aad *et al.*, [ATLAS Collab.], *Phys. Rev. D* **85**, 072004 (2012).
307. S. Chatrchyan *et al.*, [CMS Collab.], *JHEP* **1110**, 132 (2011).
308. S. Chatrchyan *et al.*, [CMS Collab.], *Phys. Rev. Lett.* **112**, 191802 (2014).
309. R. Aaij *et al.*, [LHCb Collab.], *JHEP* **1206**, 058 (2012).
310. R. Aaij *et al.*, [LHCb Collab.], *JHEP* **1508**, 039 (2015).
311. S. Chatrchyan *et al.*, [CMS Collab.], *JHEP* **1110**, 007 (2011).
312. V. Khachatryan *et al.*, [CMS Collab.], *Eur. Phys. J. C* **75**, 4 (2015).
313. G. Aad *et al.*, [ATLAS Collab.], *Phys. Lett. B* **725**, 223 (2013).
314. G. Aad *et al.*, [ATLAS Collab.], *JHEP* **1406**, 112 (2014).
315. S. Chatrchyan *et al.*, [CMS Collab.], *Phys. Rev. D* **90**, 3 (2014).
316. U. Blumenschein, arXiv:1509.04885 [hep-ex].
317. G. Aad *et al.*, [ATLAS Collab.], *JHEP* **1307**, 032 (2013).
318. J. H. Kuhn *et al.*, *JHEP* **0603**, 059 (2006).
319. V. Khachatryan *et al.*, [CMS Collab.], arXiv:1505.06520 [hep-ex].
320. M. Voutilainen, arXiv:1509.05026 [hep-ex].
321. V. Khachatryan *et al.*, [CMS Collab.], arXiv:1507.03268 [hep-ex].
322. G. Aad *et al.*, [ATLAS Collab.], *JHEP* **1301**, 086 (2013).
323. S. Chatrchyan *et al.*, [CMS Collab.], *Eur. Phys. J. C* **74**, 11 (2014).
324. K. Kröninger, A. B. Meyer, and P. Uwer, arXiv:1506.02800 [hep-ex].
325. S. Chatrchyan *et al.*, [CMS Collab.], *Phys. Lett. B* **728**, 496 (2014), *Phys. Lett. B* **728**, 526 (2014).
326. G. Aad *et al.*, [ATLAS Collab.], *Eur. Phys. J. C* **74**, 10 (2014).
327. G. Aad *et al.*, [ATLAS Collab.], *Phys. Rev. Lett.* **115**, 9 (2015).
328. S. Chatrchyan *et al.*, [CMS Collab.], *Phys. Rev. D* **89**, 9 (2014).
329. V. Khachatryan *et al.*, [CMS Collab.], arXiv:1508.07819 [hep-ex].
330. S. Bethke, *Prog. in Part. Nucl. Phys.* **58**, 351 (2007).
331. S. Bethke, *Eur. Phys. J. C* **64**, 689 (2009).
332. S. Bethke, *J. Phys. G* **26**, R27 (2000).
333. M. Schmelling, *Phys. Scripta* **51**, 676 (1995).
334. M. Beneke and M. Jamin, *JHEP* **0809**, 044 (2008).
335. K. Maltman and T. Yavin, *Phys. Rev. D* **78**, 094020 (2008).
336. S. Narison, *Phys. Lett. B* **673**, 30 (2009).
337. I. Caprini and J. Fischer, *Eur. Phys. J. C* **64**, 35 (2009).
338. A. Pich, *Prog. in Part. Nucl. Phys.* **75**, 41 (2014).
339. M. Davier *et al.*, *Eur. Phys. J. C* **74**, 2803 (2014).
340. D. Boito *et al.*, *Phys. Rev. D* **91**, 034003 (2015).
341. G. Altarelli, *PoS Corfu* **2012** (2013) 002.
342. C. McNeile *et al.*, [HPQCD Collab.], *Phys. Rev. D* **82**, 034512 (2010).
343. C.T.H. Davies *et al.*, [HPQCD Collab., UKQCD Collab., and MILC Collab.], *Phys. Rev. Lett.* **92**, 022001 (2004).
344. K. Maltman *et al.*, *Phys. Rev. D* **78**, 114504 (2008).
345. B. Chakraborty *et al.*, [HPQCD Collab.], *Phys. Rev. D* **91**, 054508 (2015).
346. S. Aoki *et al.*, [PACS-CS Collab.], *JHEP* **0910**, 053 (2009).
347. B. Blossier *et al.*, [ETM Collab.], *Phys. Rev. Lett.* **108**, 262002 (2012), *Phys. Rev. D* **89**, 014507 (2014).
348. A. Bazavov *et al.*, [BBGPSV], *Phys. Rev. D* **90**, 074038 (2014).
349. E. Shintani *et al.*, [JLQCD Collab.], *Phys. Rev. D* **82**, 074505 (2010), erratum *Phys. Rev. D* **89**, 099903 (2014).
350. S. Aoki *et al.*, *Eur. Phys. J. C* **74**, 2890 (2014).
351. C. Glasman [H1 Collab. and ZEUS Collab.], *J. Phys. Conf. Ser.* **110** 022013 (2008).
352. V. Andreev *et al.*, (H1 Collab.), *Eur. Phys. J. C* **75**, 65 (2015).
353. T. Biekötter, M. Klasen and G. Kramer, *Phys. Rev. D* **92**, 074037 (2015).
354. J. Blümlein, H. Bottcher, and A. Guffanti, *Nucl. Phys. B* **774**, 182 (2007).
355. S. Alekhin, J. Blümlein, and S. Moch, *Phys. Rev. D* **86**, 054009 (2012).
356. P. Jimenez-Delgado and E. Reya, *Phys. Rev. D* **79**, 074023 (2009).
357. A. D. Martin *et al.*, *Eur. Phys. J. C* **64**, 653 (2009).
358. L. A. Harland-Lang *et al.*, *Eur. Phys. J. C* **75**, 435 (2015).
359. R.D. Ball *et al.*, *Phys. Lett. B* **707**, 66 (2012).
360. R.S. Thorne and G. Watt, *JHEP* **1108**, 100 (2011).
361. S. Alekhin, J. Blümlein, and S. Moch, *Eur. Phys. J. C* **71**, 1723 (2011).
362. R.D. Ball *et al.*, *Phys. Lett. B* **704**, 36 (2011).
363. R.D. Ball *et al.*, *Phys. Lett. B* **723**, 330 (2013).
364. R.S. Thorne *et al.*, *PoS DIS* **2013** (2013) 042.
365. N. Brambilla *et al.*, *Phys. Rev. D* **75**, 074014 (2007).
366. G. Dissertori *et al.*, *JHEP* **0908**, 036 (2009).
367. G. Abbiendi *et al.*, *Eur. Phys. J. C* **71**, 1733 (2011).
368. S. Bethke *et al.*, [JADE Collab.], *Eur. Phys. J. C* **64**, 351 (2009).
369. G. Dissertori *et al.*, *Phys. Rev. Lett.* **104**, 072002 (2010).
370. J. Schieck *et al.*, *Eur. Phys. J. C* **73**, 2332 (2013).
371. R.A. Davison and B.R. Webber, *Eur. Phys. J. C* **59**, 13 (2009).
372. R. Abbate *et al.*, *Phys. Rev. D* **83**, 074021 (2011).
373. T. Gehrmann *et al.*, *Eur. Phys. J. C* **73**, 2265 (2013).
374. A.H. Hoang *et al.*, *Phys. Rev. D* **91**, 094018 (2015).
375. R. Frederix *et al.*, *JHEP* **1011**, 050 (2010).
376. P. Bolzoni, B.A. Kniehl, and A.V. Kotikov, *Nucl. Phys. B* **875**, 18 (2013).
377. M. Abazov *et al.*, [D0 Collab.], *Phys. Rev. D* **80**, 111107 (2009).
378. M. Abazov *et al.*, [D0 Collab.], *Phys. Lett. B* **718**, 56 (2012).
379. G. Aad *et al.*, [ATLAS Collab.], *Phys. Rev. D* **86**, 014022 (2012).
380. B. Malaescu and P. Starovoitov, *Eur. Phys. J. C* **72**, 2041 (2012).
381. S. Chatrchyan *et al.*, [CMS Collab.], *Eur. Phys. J. C* **73**, 2604 (2013).
382. V. Khachatryan *et al.*, [CMS Collab.], *Eur. Phys. J. C* **75**, 288 (2015).
383. V. Khachatryan *et al.*, [CMS Collab.], *Eur. Phys. J. C* **75**, 186 (2015).
384. M. Baak *et al.*, [Gitter group], *Eur. Phys. J. C* **74**, 304660 (2014).



## 10. ELECTROWEAK MODEL AND CONSTRAINTS ON NEW PHYSICS

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### 10.1. Introduction

The standard model of the electroweak interactions (SM) [1] is based on the gauge group  $SU(2) \times U(1)$ , with gauge bosons  $W_\mu^i$ ,  $i = 1, 2, 3$ , and  $B_\mu$  for the  $SU(2)$  and  $U(1)$  factors, respectively, and the corresponding gauge coupling constants  $g$  and  $g'$ . The left-handed fermion fields of the  $i^{\text{th}}$  fermion family transform as doublets  $\Psi_i = \begin{pmatrix} \nu_i \\ \ell_i \end{pmatrix}$  and  $\begin{pmatrix} u_i \\ d_i \end{pmatrix}$  under  $SU(2)$ , where  $d_i \equiv \sum_j V_{ij} d_j$ , and  $V$  is the Cabibbo-Kobayashi-Maskawa mixing matrix. [Constraints on  $V$  and tests of universality are discussed in Ref. 2 and in the Section on “The CKM Quark-Mixing Matrix”. The extension of the formalism to allow an analogous leptonic mixing matrix is discussed in the Section on “Neutrino Mass, Mixing, and Oscillations”.] The right-handed fields are  $SU(2)$  singlets. From Higgs and electroweak precision data it is known that there are precisely three sequential fermion families.

A complex scalar Higgs doublet,  $\phi \equiv \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$ , is added to the model for mass generation through spontaneous symmetry breaking with potential\* given by,

$$V(\phi) = \mu^2 \phi^\dagger \phi + \frac{\lambda^2}{2} (\phi^\dagger \phi)^2. \quad (10.1)$$

For  $\mu^2$  negative,  $\phi$  develops a vacuum expectation value,  $v/\sqrt{2} = \mu/\lambda$ , where  $v \approx 246$  GeV, breaking part of the electroweak (EW) gauge symmetry, after which only one neutral Higgs scalar,  $H$ , remains in the physical particle spectrum. In non-minimal models there are additional charged and neutral scalar Higgs particles [3].

After the symmetry breaking the Lagrangian for the fermion fields,  $\psi_i$ , is

$$\begin{aligned} \mathcal{L}_F = & \sum_i \bar{\psi}_i \left( i \not{\partial} - m_i - \frac{m_i H}{v} \right) \psi_i \\ & - \frac{g}{2\sqrt{2}} \sum_i \bar{\psi}_i \gamma^\mu (1 - \gamma^5) (T^+ W_\mu^+ + T^- W_\mu^-) \psi_i \\ & - e \sum_i Q_i \bar{\psi}_i \gamma^\mu \psi_i A_\mu \\ & - \frac{g}{2 \cos \theta_W} \sum_i \bar{\psi}_i \gamma^\mu (g_V^i - g_A^i \gamma^5) \psi_i Z_\mu. \end{aligned} \quad (10.2)$$

Here  $\theta_W \equiv \tan^{-1}(g'/g)$  is the weak angle;  $e = g \sin \theta_W$  is the positron electric charge; and  $A \equiv B \cos \theta_W + W^3 \sin \theta_W$  is the photon field ( $\gamma$ ).  $W^\pm \equiv (W^1 \mp iW^2)/\sqrt{2}$  and  $Z \equiv -B \sin \theta_W + W^3 \cos \theta_W$  are the charged and neutral weak boson fields, respectively. The Yukawa coupling of  $H$  to  $\psi_i$  in the first term in  $\mathcal{L}_F$ , which is flavor diagonal in the minimal model, is  $gm_i/2M_W$ . The boson masses in the EW sector are given (at tree level, *i.e.*, to lowest order in perturbation theory) by,

$$M_H = \lambda v, \quad (10.3a)$$

$$M_W = \frac{1}{2} g v = \frac{e v}{2 \sin \theta_W}, \quad (10.3b)$$

$$M_Z = \frac{1}{2} \sqrt{g^2 + g'^2} v = \frac{e v}{2 \sin \theta_W \cos \theta_W} = \frac{M_W}{\cos \theta_W}, \quad (10.3c)$$

$$M_\gamma = 0. \quad (10.3d)$$

\* There is no generally accepted convention to write the quartic term. Our numerical coefficient simplifies Eq. (10.3a) below and the squared coupling preserves the relation between the number of external legs and the power counting of couplings at a given loop order. This structure also naturally emerges from physics beyond the SM, such as supersymmetry.

The second term in  $\mathcal{L}_F$  represents the charged-current weak interaction [4–7], where  $T^+$  and  $T^-$  are the weak isospin raising and lowering operators. For example, the coupling of a  $W$  to an electron and a neutrino is

$$-\frac{e}{2\sqrt{2} \sin \theta_W} \left[ W_\mu^- \bar{\nu} \gamma^\mu (1 - \gamma^5) \nu + W_\mu^+ \bar{\nu} \gamma^\mu (1 - \gamma^5) e \right]. \quad (10.4)$$

For momenta small compared to  $M_W$ , this term gives rise to the effective four-fermion interaction with the Fermi constant given by  $G_F/\sqrt{2} = 1/2v^2 = g^2/8M_W^2$ .  $CP$  violation is incorporated into the EW model by a single observable phase in  $V_{ij}$ .

The third term in  $\mathcal{L}_F$  describes electromagnetic interactions (QED) [8,9], and the last is the weak neutral-current interaction [5–7]. The vector and axial-vector couplings are

$$g_V^i \equiv t_{3L}(i) - 2Q_i \sin^2 \theta_W, \quad (10.5a)$$

$$g_A^i \equiv t_{3L}(i), \quad (10.5b)$$

where  $t_{3L}(i)$  is the weak isospin of fermion  $i$  ( $+1/2$  for  $u_i$  and  $\nu_i$ ;  $-1/2$  for  $d_i$  and  $e_i$ ) and  $Q_i$  is the charge of  $\psi_i$  in units of  $e$ .

The first term in Eq. (10.2) also gives rise to fermion masses, and in the presence of right-handed neutrinos to Dirac neutrino masses. The possibility of Majorana masses is discussed in the Section on “Neutrino Mass, Mixing, and Oscillations”.

### 10.2. Renormalization and radiative corrections

In addition to the Higgs boson mass,  $M_H$ , the fermion masses and mixings, and the strong coupling constant,  $\alpha_s$ , the SM has three parameters. The set with the smallest experimental errors contains the  $Z$  mass\*\*, the Fermi constant, and the fine structure constant, which will be discussed in turn (if not stated otherwise, the numerical values quoted in Sec. 10.2–10.5 correspond to the main fit result in Table 10.6):

The  $Z$  boson mass,  $M_Z = 91.1876 \pm 0.0021$  GeV, has been determined from the  $Z$  lineshape scan at LEP 1 [10]. This value of  $M_Z$  corresponds to a definition based on a Breit-Wigner shape with an energy-dependent width (see the Section on “The  $Z$  Boson” in the Gauge and Higgs Boson Particle Listings of this *Review*).

The Fermi constant,  $G_F = 1.1663787(6) \times 10^{-5}$  GeV $^{-2}$ , is derived from the muon lifetime formula\*\*\*,

$$\frac{\hbar}{\tau_\mu} = \frac{G_F^2 m_\mu^5}{192\pi^3} F(\rho) \left[ 1 + H_1(\rho) \frac{\hat{\alpha}(m_\mu)}{\pi} + H_2(\rho) \frac{\hat{\alpha}^2(m_\mu)}{\pi^2} \right], \quad (10.6)$$

where  $\rho = m_e^2/m_\mu^2$ , and where

$$F(\rho) = 1 - 8\rho + 8\rho^3 - \rho^4 - 12\rho^2 \ln \rho = 0.99981295, \quad (10.7a)$$

$$\begin{aligned} H_1(\rho) = & \frac{25}{8} - \frac{\pi^2}{2} - \left( 9 + 4\pi^2 + 12 \ln \rho \right) \rho \\ & + 16\pi^2 \rho^{3/2} + \mathcal{O}(\rho^2) = -1.80793, \end{aligned} \quad (10.7b)$$

$$\begin{aligned} H_2(\rho) = & \frac{156815}{5184} - \frac{518}{81} \pi^2 - \frac{895}{36} \zeta(3) + \frac{67}{720} \pi^4 + \frac{53}{6} \pi^2 \ln 2 \\ & - (0.042 \pm 0.002)_{\text{had}} - \frac{5}{4} \pi^2 \sqrt{\rho} + \mathcal{O}(\rho) = 6.64, \end{aligned} \quad (10.7c)$$

$$\hat{\alpha}(m_\mu)^{-1} = \alpha^{-1} + \frac{1}{3\pi} \ln \rho + \mathcal{O}(\alpha) = 135.901 \quad (10.7d)$$

\*\* We emphasize that in the fits described in Sec. 10.6 and Sec. 10.7 the values of the SM parameters are affected by all observables that depend on them. This is of no practical consequence for  $\alpha$  and  $G_F$ , however, since they are very precisely known.

\*\*\* In the spirit of the Fermi theory, we incorporated the small propagator correction,  $3/5 m_\mu^2/M_W^2$ , into  $\Delta r$  (see below). This is also the convention adopted by the MuLan collaboration [11]. While this breaks with historical consistency, the numerical difference was negligible in the past.

$H_1$  and  $H_2$  capture the QED corrections within the Fermi model. The results for  $\rho = 0$  have been obtained in Refs. 12 and 13, respectively, where the term in parentheses is from the hadronic vacuum polarization [13]. The mass corrections to  $H_1$  have been known for some time [14], while those to  $H_2$  are more recent [15]. Notice the term linear in  $m_e$  whose appearance was unforeseen and can be traced to the use of the muon pole mass in the prefactor [15]. The remaining uncertainty in  $G_F$  is experimental and has recently been reduced by an order of magnitude by the MuLan collaboration [11] at the PSI.

The experimental determination of the fine structure constant,  $\alpha = 1/137.035999139(31)$ , is currently dominated by the  $e^\pm$  anomalous magnetic moment [16]. In most EW renormalization schemes, it is convenient to define a running  $\alpha$  dependent on the energy scale of the process, with  $\alpha^{-1} \sim 137$  appropriate at very low energy, *i.e.* close to the Thomson limit. (The running has also been observed [17] directly.) For scales above a few hundred MeV this introduces an uncertainty due to the low energy hadronic contribution to vacuum polarization. In the modified minimal subtraction ( $\overline{\text{MS}}$ ) scheme [18] (used for this *Review*), and with  $\alpha_s(M_Z) = 0.1182 \pm 0.0016$  we have  $\hat{\alpha}(m_\tau)^{-1} = 133.471 \pm 0.016$  and  $\hat{\alpha}(M_Z)^{-1} = 127.950 \pm 0.017$ . (In this Section we denote quantities defined in the modified minimal subtraction ( $\overline{\text{MS}}$ ) scheme by a caret; the exception is the strong coupling constant,  $\alpha_s$ , which will always correspond to the  $\overline{\text{MS}}$  definition and where the caret will be dropped.) The latter corresponds to a quark sector contribution (without the top) to the conventional (on-shell) QED coupling,  $\alpha(M_Z) = \frac{\alpha}{1 - \Delta\alpha(M_Z)}$ , of

$\Delta\alpha_{\text{had}}^{(5)}(M_Z) = 0.02764 \pm 0.00013$ . These values are updated from Ref. 19 with  $\Delta\alpha_{\text{had}}^{(5)}(M_Z)$  moved downwards and its uncertainty reduced (partly due to a more precise charm quark mass). Its correlation with the  $\mu^\pm$  anomalous magnetic moment (see Sec. 10.5), as well as the non-linear  $\alpha_s$  dependence of  $\hat{\alpha}(M_Z)$  and the resulting correlation with the input variable  $\alpha_s$ , are fully taken into account in the fits. This is done by using as actual input (fit constraint) instead of  $\Delta\alpha_{\text{had}}^{(5)}(M_Z)$  the low energy contribution by the three light quarks,  $\Delta\alpha_{\text{had}}^{(3)}(2.0 \text{ GeV}) = (58.04 \pm 1.10) \times 10^{-4}$  [20], and by calculating the perturbative and heavy quark contributions to  $\hat{\alpha}(M_Z)$  in each call of the fits according to Ref. 19. Part of the uncertainty ( $\pm 0.92 \times 10^{-4}$ ) is from  $e^+e^-$  annihilation data below 1.8 GeV and  $\tau$  decay data (including uncertainties from isospin breaking effects), but uncalculated higher order perturbative ( $\pm 0.41 \times 10^{-4}$ ) and non-perturbative ( $\pm 0.44 \times 10^{-4}$ ) QCD corrections and the  $\overline{\text{MS}}$  quark mass values (see below) also contribute. Various evaluations of  $\Delta\alpha_{\text{had}}^{(5)}$  are summarized in Table 10.1 where the relation<sup>†</sup> between the  $\overline{\text{MS}}$  and on-shell definitions (obtained using Ref. 23) is given by,

$$\begin{aligned} \Delta\hat{\alpha}(M_Z) - \Delta\alpha(M_Z) = & \frac{\alpha}{\pi} \left[ \left( \frac{100}{27} - \frac{1}{6} - \frac{7}{4} \ln \frac{M_Z^2}{M_W^2} \right) + \frac{\alpha_s(M_Z)}{\pi} \left( \frac{605}{108} - \frac{44}{9} \zeta(3) \right) \right. \\ & \left. + \frac{\alpha_s^2(M_Z)}{\pi^2} \left( \frac{976481}{23328} - \frac{253}{36} \zeta(2) - \frac{781}{18} \zeta(3) + \frac{275}{27} \zeta(5) \right) \right] = \\ & = 0.007127(2), \end{aligned} \quad (10.8)$$

and where the first entry of the lowest order term is from fermions and the other two are from  $W^\pm$  loops, which are usually excluded from the on-shell definition. The most recent results typically assume the validity of perturbative QCD (PQCD) at scales of 1.8 GeV or above, and are in reasonable agreement with each other. In regions where PQCD is not trusted, one can use  $e^+e^- \rightarrow \text{hadrons}$  cross-section data and  $\tau$  decay spectral functions [31], where the latter derive from OPAL [34], CLEO [35], ALEPH [36], and Belle [37]. The dominant  $e^+e^- \rightarrow \pi^+\pi^-$  cross-section was measured with the CMD-2 [38] and

<sup>†</sup> In practice,  $\alpha(M_Z)$  is directly evaluated in the  $\overline{\text{MS}}$  scheme using the FORTRAN package GAPP [21], including the QED contributions of both leptons and quarks. The leptonic three-loop contribution in the on-shell scheme has been obtained in Ref. 22.

Reference	Result	Comment
Geshkenbein, Morgunov [24]	$0.02780 \pm 0.00006$	$\mathcal{O}(\alpha_s)$ resonance model
Swartz [25]	$0.02754 \pm 0.00046$	use of fitting function
Krasnikov, Rodenberg [26]	$0.02737 \pm 0.00039$	PQCD for $\sqrt{s} > 2.3 \text{ GeV}$
Kühn & Steinhauser [27]	$0.02778 \pm 0.00016$	full $\mathcal{O}(\alpha_s^2)$ for $\sqrt{s} > 1.8 \text{ GeV}$
Erler [19]	$0.02779 \pm 0.00020$	conv. from $\overline{\text{MS}}$ scheme
Groote <i>et al.</i> [28]	$0.02787 \pm 0.00032$	use of QCD sum rules
Martin <i>et al.</i> [29]	$0.02741 \pm 0.00019$	incl. new BES data
de Troconiz, Yndurain [30]	$0.02754 \pm 0.00010$	PQCD for $s > 2 \text{ GeV}^2$
Davier <i>et al.</i> [31]	$0.02762 \pm 0.00011$	incl. $\tau$ decay data PQCD for $\sqrt{s} > 1.8 \text{ GeV}$
Burkhardt, Pietrzyk [32]	$0.02750 \pm 0.00033$	incl. BES/BABAR data, PQCD for $\sqrt{s} > 12 \text{ GeV}$
Hagiwara <i>et al.</i> [33]	$0.02764 \pm 0.00014$	incl. new $e^+e^-$ data, PQCD for $\sqrt{s} =$ $= 2.6\text{--}3.7, > 11.1 \text{ GeV}$
Jegerlehner [20]	$0.02766 \pm 0.00018$	incl. $\gamma$ - $\rho$ mixing corrected $\tau$ data, PQCD: $\sqrt{s} = 5.2\text{--}9.46, > 13 \text{ GeV}$

**Table 10.1:** Evaluations of the on-shell  $\Delta\alpha_{\text{had}}^{(5)}(M_Z)$  by different groups (for a more complete list of evaluations see the 2012 edition of this *Review*). For better comparison we adjusted central values and errors to correspond to a common and fixed value of  $\alpha_s(M_Z) = 0.120$ . References quoting results without the top quark decoupled are converted to the five flavor definition. Ref. [28] uses  $\Lambda_{\text{QCD}} = 380 \pm 60 \text{ MeV}$ ; for the conversion we assumed  $\alpha_s(M_Z) = 0.118 \pm 0.003$ .

SND [39] detectors at the VEPP-2M  $e^+e^-$  collider at Novosibirsk. As an alternative to cross-section scans, one can use the high statistics radiative return events at  $e^+e^-$  accelerators operating at resonances such as the  $\Phi$  or the  $\Upsilon(4S)$ . The method [40] is systematics limited but dominates over the Novosibirsk data throughout. The BaBar collaboration [41] studied multi-hadron events radiatively returned from the  $\Upsilon(4S)$ , reconstructing the radiated photon and normalizing to  $\mu^\pm\gamma$  final states. Their result is higher compared to VEPP-2M, while the shape and smaller overall cross-section from the  $\pi^+\pi^-$  radiative return results from the  $\Phi$  obtained by the KLOE collaboration [42] differ significantly from what is observed by BaBar. The discrepancy originates from the kinematic region  $\sqrt{s} \gtrsim 0.6 \text{ GeV}$ , and is most pronounced for  $\sqrt{s} \gtrsim 0.85 \text{ GeV}$ . All measurements including older data [43] and multi-hadron final states (there are also discrepancies in the  $e^+e^- \rightarrow 2\pi^+2\pi^-$  channel [31]) are accounted for and corrections have been applied for missing channels. Further improvement of this dominant theoretical uncertainty in the interpretation of precision data will require better measurements of the cross-section for  $e^+e^- \rightarrow \text{hadrons}$  below the charmonium resonances including multi-pion and other final states. To improve the precisions in  $\hat{m}_c(\hat{m}_c)$  and  $\hat{m}_b(\hat{m}_b)$  it would help to remeasure the threshold regions of the heavy quarks as well as the electronic decay widths of the narrow  $c\bar{c}$  and  $b\bar{b}$  resonances.

Further free parameters entering into Eq. (10.2) are the quark and lepton masses, where  $m_i$  is the mass of the  $i^{\text{th}}$  fermion  $\psi_i$ . For the light quarks, as described in the note on “Quark Masses” in the Quark Listings,  $\hat{m}_u = 2.3^{+0.7}_{-0.5} \text{ MeV}$ ,  $\hat{m}_d = 4.8^{+0.5}_{-0.3} \text{ MeV}$ , and  $\hat{m}_s = 95 \pm 5 \text{ MeV}$ . These are running  $\overline{\text{MS}}$  masses evaluated at the scale  $\mu = 2 \text{ GeV}$ . For the heavier quarks we use QCD sum rule [44] constraints [45] and recalculate their masses in each call of our fits to account for their direct  $\alpha_s$  dependence. We find<sup>¶</sup>,

<sup>¶</sup> Other authors [46] advocate to evaluate and quote  $\hat{m}_c(\mu = 3 \text{ GeV})$  instead. We use  $\hat{m}_c(\mu = \hat{m}_c)$  because in the global analysis it is convenient to nullify any explicitly  $m_c$  dependent logarithms. Note also that



$\hat{m}_c(\mu = \hat{m}_c) = 1.265^{+0.030}_{-0.038}$  GeV and  $\hat{m}_b(\mu = \hat{m}_b) = 4.199 \pm 0.023$  GeV, with a correlation of 23%.

There are two recent combinations of measurements of the top quark “pole” mass (the quotation marks are a reminder that the experiments do not strictly measure the pole mass and that quarks do not form asymptotic states), all of them utilizing kinematic reconstruction. The most recent result, which we will use as our default input value,  $m_t = 173.34 \pm 0.37_{\text{stat.}} \pm 0.52_{\text{syst.}}$  GeV [48], is internal to the Tevatron and combines 12 individual CDF and DØ measurements including Run I and other less precise determinations. The other one,  $m_t = 173.34 \pm 0.27_{\text{stat.}} \pm 0.71_{\text{syst.}}$  GeV [49], is a Tevatron/LHC combination based on a selection of 11 individual results. The above averages differ slightly from the value,  $m_t = 173.21 \pm 0.51_{\text{stat.}} \pm 0.71_{\text{syst.}}$  GeV, which appears in the top quark Listings in this *Review* and which is based exclusively on published Tevatron results. We are working, however, with  $\overline{\text{MS}}$  masses in all expressions to minimize theoretical uncertainties. Such a short distance mass definition (unlike the pole mass) is free from non-perturbative and renormalon [50] uncertainties. We therefore convert to the top quark  $\overline{\text{MS}}$  mass using the three-loop formula [51]. (Very recently, the four-loop result has been obtained in Ref. 52.) This introduces an additional uncertainty which we estimate to 0.5 GeV (the size of the three-loop term) and add in quadrature to the experimental pole mass error. This is convenient because we use the pole mass as an external constraint while fitting to the  $\overline{\text{MS}}$  mass. We are assuming that the kinematic mass extracted from the collider events corresponds within this uncertainty to the pole mass. In summary, we will use the fit constraint,

$$m_t = 173.34 \pm 0.64_{\text{exp.}} \pm 0.5_{\text{QCD}} \text{ GeV} = 173.34 \pm 0.81 \text{ GeV}. \quad (10.9)$$

While there seems to be perfect agreement between all these averages, we observe a  $2.6 \sigma$  deviation (or more in case of correlated systematics) between the two most precise determinations,  $174.98 \pm 0.76$  GeV [53] (by the DØ Collaboration) and  $172.22 \pm 0.73$  GeV [54] (by the CMS Collaboration), both from the lepton + jets channels [55]. For more details, see the Section on “The Top Quark” and the Quarks Listings in this *Review*.

The observables  $\sin^2 \theta_W$  and  $M_W$  can be calculated from  $M_Z$ ,  $\hat{\alpha}(M_Z)$ , and  $G_F$ , when values for  $m_t$  and  $M_H$  are given, or conversely,  $M_H$  can be constrained by  $\sin^2 \theta_W$  and  $M_W$ . The value of  $\sin^2 \theta_W$  is extracted from neutral-current processes (see Sec. 10.3) and  $Z$  pole observables (see Sec. 10.4) and depends on the renormalization prescription. There are a number of popular schemes [56–62] leading to values which differ by small factors depending on  $m_t$  and  $M_H$ . The notation for these schemes is shown in Table 10.2.

**Table 10.2:** Notations used to indicate the various schemes discussed in the text. Each definition of  $\sin^2 \theta_W$  leads to values that differ by small factors depending on  $m_t$  and  $M_H$ . Numerical values and the uncertainties induced by the imperfectly known SM parameters are also given for illustration.

Scheme	Notation	Value	Parametric uncertainty
On-shell	$s_W^2$	0.22336	$\pm 0.00010$
$\overline{\text{MS}}$	$\hat{s}_Z^2$	0.23129	$\pm 0.00005$
$\overline{\text{MS}}$ ND	$\hat{s}_{\text{ND}}^2$	0.23148	$\pm 0.00005$
$\overline{\text{MS}}$	$\hat{s}_0^2$	0.23865	$\pm 0.00008$
Effective angle	$\hat{s}_\ell^2$	0.23152	$\pm 0.00005$

our uncertainty for  $m_c$  (and to a lesser degree for  $m_b$ ) is larger than in Refs. 46 and 47, for example. The reason is that we determine the continuum contribution for charm pair production using only resonance data and theoretical consistency across various sum rule moments, and then use any difference to the experimental continuum data as an additional uncertainty. We also include an uncertainty for the condensate terms which grows rapidly for higher moments in the sum rule analysis.

- (i) The on-shell scheme [56] promotes the tree-level formula  $\sin^2 \theta_W = 1 - M_W^2/M_Z^2$  to a definition of the renormalized  $\sin^2 \theta_W$  to all orders in perturbation theory, i.e.,  $\sin^2 \theta_W \rightarrow s_W^2 \equiv 1 - M_W^2/M_Z^2$ :

$$M_W = \frac{A_0}{s_W(1 - \Delta r)^{1/2}}, \quad M_Z = \frac{M_W}{c_W}, \quad (10.10)$$

where  $c_W \equiv \cos \theta_W$ ,  $A_0 = (\pi\alpha/\sqrt{2}G_F)^{1/2} = 37.28039(1)$  GeV, and  $\Delta r$  includes the radiative corrections relating  $\alpha$ ,  $\alpha(M_Z)$ ,  $G_F$ ,  $M_W$ , and  $M_Z$ . One finds  $\Delta r \sim \Delta r_0 - \rho_t/\tan^2 \theta_W$ , where  $\Delta r_0 = 1 - \alpha/\hat{\alpha}(M_Z) = 0.06630(13)$  is due to the running of  $\alpha$ , and  $\rho_t = 3G_F m_t^2/8\sqrt{2}\pi^2 = 0.00940 (m_t/173.34 \text{ GeV})^2$  represents the dominant (quadratic)  $m_t$  dependence. There are additional contributions to  $\Delta r$  from bosonic loops, including those which depend logarithmically on  $M_H$  and higher-order corrections<sup>§§</sup>. One has  $\Delta r = 0.03648 \mp 0.00028 \pm 0.00013$ , where the first uncertainty is from  $m_t$  and the second is from  $\alpha(M_Z)$ . Thus the value of  $s_W^2$  extracted from  $M_Z$  includes an uncertainty ( $\mp 0.00009$ ) from the currently allowed range of  $m_t$ . This scheme is simple conceptually. However, the relatively large ( $\sim 3\%$ ) correction from  $\rho_t$  causes large spurious contributions in higher orders.

$s_W^2$  depends not only on the gauge couplings but also on the spontaneous-symmetry breaking, and it is awkward in the presence of any extension of the SM which perturbs the value of  $M_Z$  (or  $M_W$ ). Other definitions are motivated by the tree-level coupling constant definition  $\theta_W = \tan^{-1}(g'/g)$ :

- (ii) In particular, the modified minimal subtraction ( $\overline{\text{MS}}$ ) scheme introduces the quantity  $\sin^2 \hat{\theta}_W(\mu) \equiv \hat{g}^2(\mu)/[\hat{g}^2(\mu) + \hat{g}'^2(\mu)]$ , where the couplings  $\hat{g}$  and  $\hat{g}'$  are defined by modified minimal subtraction and the scale  $\mu$  is conveniently chosen to be  $M_Z$  for many EW processes. The value of  $\hat{s}_Z^2 = \sin^2 \hat{\theta}_W(M_Z)$  extracted from  $M_Z$  is less sensitive than  $s_W^2$  to  $m_t$  (by a factor of  $\tan^2 \theta_W$ ), and is less sensitive to most types of new physics. It is also very useful for comparing with the predictions of grand unification. There are actually several variant definitions of  $\sin^2 \hat{\theta}_W(M_Z)$ , differing according to whether or how finite  $\alpha \ln(m_t/M_Z)$  terms are decoupled (subtracted from the couplings). One cannot entirely decouple the  $\alpha \ln(m_t/M_Z)$  terms from all EW quantities because  $m_t \gg m_b$  breaks SU(2) symmetry. The scheme that will be adopted here decouples the  $\alpha \ln(m_t/M_Z)$  terms from the  $\gamma$ - $Z$  mixing [18,57], essentially eliminating any  $\ln(m_t/M_Z)$  dependence in the formulae for asymmetries at the  $Z$  pole when written in terms of  $\hat{s}_Z^2$ . (A similar definition is used for  $\hat{\alpha}$ .) The on-shell and  $\overline{\text{MS}}$  definitions are related by

$$\hat{s}_Z^2 = c(m_t, M_H) s_W^2 = (1.0355 \pm 0.0003) s_W^2. \quad (10.11)$$

The quadratic  $m_t$  dependence is given by  $c \sim 1 + \rho_t/\tan^2 \theta_W$ . The expressions for  $M_W$  and  $M_Z$  in the  $\overline{\text{MS}}$  scheme are

$$M_W = \frac{A_0}{\hat{s}_Z(1 - \Delta \hat{r}_W)^{1/2}}, \quad M_Z = \frac{M_W}{\hat{\rho}^{1/2} \hat{c}_Z}, \quad (10.12)$$

and one predicts  $\Delta \hat{r}_W = 0.06952 \pm 0.00013$ .  $\Delta \hat{r}_W$  has no quadratic  $m_t$  dependence, because shifts in  $M_W$  are absorbed into the observed  $G_F$ , so that the error in  $\Delta \hat{r}_W$  is almost entirely due to  $\Delta r_0 = 1 - \alpha/\hat{\alpha}(M_Z)$ . The quadratic  $m_t$  dependence has been shifted into  $\hat{\rho} \sim 1 + \rho_t$ , where including bosonic loops,  $\hat{\rho} = 1.01032 \pm 0.00009$ .

- (iii) A variant  $\overline{\text{MS}}$  quantity  $\hat{s}_{\text{ND}}^2$  (used in the 1992 edition of this *Review*) does not decouple the  $\alpha \ln(m_t/M_Z)$  terms [58]. It is related to  $\hat{s}_Z^2$  by

$$\hat{s}_Z^2 = \hat{s}_{\text{ND}}^2 / \left(1 + \frac{\hat{\alpha}}{\pi} d\right), \quad (10.13a)$$

$$d = \frac{1}{3} \left( \frac{1}{\hat{s}_Z^2} - \frac{8}{3} \right) \left[ \left(1 + \frac{\alpha_s}{\pi}\right) \ln \frac{m_t}{M_Z} - \frac{15\alpha_s}{8\pi} \right], \quad (10.13b)$$

<sup>§§</sup> All explicit numbers quoted here and below include the two- and three-loop corrections described near the end of Sec. 10.2.

Thus,  $\hat{s}_Z^2 - \hat{s}_{\text{ND}}^2 \approx -0.0002$ .

(iv) Some of the low-energy experiments discussed in the next section are sensitive to the weak mixing angle at almost vanishing momentum transfer (for a review, see Ref. 59). Thus, Table 10.2 also includes  $\hat{s}_0^2 \equiv \sin^2 \theta_W(0)$ .

(v) Yet another definition, the effective angle [60–62]  $\bar{s}_f^2 = \sin^2 \theta_{\text{eff}}^f$  for the  $Z$  vector coupling to fermion  $f$ , is based on  $Z$  pole observables and described in Sec. 10.4.

Experiments are at such level of precision that complete one-loop, dominant two-loop, and partial three and four-loop radiative corrections must be applied. For neutral-current and  $Z$  pole processes, these corrections are conveniently divided into two classes:

1. QED diagrams involving the emission of real photons or the exchange of virtual photons in loops, but not including vacuum polarization diagrams. These graphs often yield finite and gauge-invariant contributions to observable processes. However, they are dependent on energies, experimental cuts, *etc.*, and must be calculated individually for each experiment.
2. EW corrections, including  $\gamma\gamma$ ,  $\gamma Z$ ,  $ZZ$ , and  $WW$  vacuum polarization diagrams, as well as vertex corrections, box graphs, *etc.*, involving virtual  $W$  and  $Z$  bosons. The one-loop corrections are included for all processes, and many two-loop corrections are also important. In particular, two-loop corrections involving the top quark modify  $\rho_t$  in  $\hat{\rho}$ ,  $\Delta r$ , and elsewhere by

$$\rho_t \rightarrow \rho_t[1 + R(M_H, m_t)\rho_t/3]. \quad (10.14)$$

$R(M_H, m_t)$  can be described as an expansion in  $M_Z^2/m_t^2$ , for which the leading  $m_t^4/M_Z^4$  [63] and next-to-leading  $m_t^2/M_Z^2$  [64] terms are known. The complete two-loop calculation of  $\Delta r$  (without further approximation) has been performed in Refs. 65 and 66 for fermionic and purely bosonic diagrams, respectively. Similarly, the EW two-loop calculation for the relation between  $\bar{s}_\ell^2$  and  $s_W^2$  is complete [67,68]. Very recently, Ref. 69 obtained the  $\overline{\text{MS}}$  quantities  $\Delta r_W$  and  $\hat{\rho}$  to two-loop accuracy, confirming the prediction of  $M_W$  in the on-shell scheme from Refs. 66 and 70 within about 4 MeV.

Mixed QCD-EW contributions to gauge boson self-energies of order  $\alpha\alpha_s m_t^2$  [71],  $\alpha\alpha_s^2 m_t^2$  [72], and  $\alpha\alpha_s^3 m_t^2$  [73] increase the predicted value of  $m_t$  by 6%. This is, however, almost entirely an artifact of using the pole mass definition for  $m_t$ . The equivalent corrections when using the  $\overline{\text{MS}}$  definition  $\hat{m}_t(\hat{m}_t)$  increase  $m_t$  by less than 0.5%. The subleading  $\alpha\alpha_s$  corrections [74] are also included. Further three-loop corrections of order  $\alpha\alpha_s^2$  [75,76],  $\alpha^3 m_t^6$ , and  $\alpha^2 \alpha_s m_t^4$  [77], are rather small. The same is true for  $\alpha^3 M_H^4$  [78] corrections unless  $M_H$  approaches 1 TeV.

The theoretical uncertainty from unknown higher-order corrections is estimated to amount to 4 MeV for the prediction of  $M_W$  [70] and  $4.5 \times 10^{-5}$  for  $\bar{s}_\ell^2$  [79].

Throughout this *Review* we utilize EW radiative corrections from the program GAPP [21], which works entirely in the  $\overline{\text{MS}}$  scheme, and which is independent of the package ZFITTER [62].

### 10.3. Low energy electroweak observables

In the following we discuss EW precision observables obtained at low momentum transfers [6], *i.e.*  $Q^2 \ll M_Z^2$ . It is convenient to write the four-fermion interactions relevant to  $\nu$ -hadron,  $\nu$ - $e$ , as well as parity violating  $e$ -hadron and  $e$ - $e$  neutral-current processes in a form that is valid in an arbitrary gauge theory (assuming massless

left-handed neutrinos). One has<sup>★</sup>

$$-\mathcal{L}^{\nu e} = \frac{G_F}{\sqrt{2}} \bar{\nu} \gamma_\mu (1 - \gamma^5) \nu \bar{e} \gamma^\mu (g_{LV}^{\nu e} - g_{LA}^{\nu e} \gamma^5) e, \quad (10.15)$$

$$-\mathcal{L}^{\nu h} = \frac{G_F}{\sqrt{2}} \bar{\nu} \gamma_\mu (1 - \gamma^5) \nu \sum_q [g_{LL}^{\nu q} \bar{q} \gamma^\mu (1 - \gamma^5) q + g_{LR}^{\nu q} \bar{q} \gamma^\mu (1 + \gamma^5) q], \quad (10.16)$$

$$-\mathcal{L}^{ee} = -\frac{G_F}{\sqrt{2}} g_{AV}^{ee} \bar{e} \gamma_\mu \gamma^5 e \bar{e} \gamma^\mu e, \quad (10.17)$$

$$-\mathcal{L}^{eh} = -\frac{G_F}{\sqrt{2}} \sum_q [g_{AV}^{eq} \bar{e} \gamma_\mu \gamma^5 e \bar{q} \gamma^\mu q + g_{VA}^{eq} \bar{e} \gamma_\mu e \bar{q} \gamma^\mu \gamma^5 q], \quad (10.18)$$

where one must include the charged-current contribution for  $\nu_e$ - $e$  and  $\bar{\nu}_e$ - $e$  and the parity-conserving QED contribution for electron scattering.

**Table 10.3:** SM tree level expressions for the neutral-current parameters for  $\nu$ -hadron,  $\nu$ - $e$ , and  $e$ - $e$ -scattering processes. To obtain the SM values in the last column, the tree level expressions have to be multiplied by the low-energy neutral-current  $\rho$  parameter,  $\rho_{\text{NC}} = 1.00066$ , and further vertex and box corrections need to be added as detailed in Ref. 80. The dominant  $m_t$  dependence is again given by  $\rho_{\text{NC}} \sim 1 + \rho_t$ .

Quantity	SM tree level	SM value
$g_{LV}^{\nu e}$	$-\frac{1}{2} + 2 \hat{s}_0^2$	-0.0396
$g_{LA}^{\nu e}$	$-\frac{1}{2}$	-0.5064
$g_{LL}^{\nu u}$	$\frac{1}{2} - \frac{2}{3} \hat{s}_0^2$	0.3457
$g_{LL}^{\nu d}$	$-\frac{1}{2} + \frac{1}{3} \hat{s}_0^2$	-0.4288
$g_{LR}^{\nu u}$	$-\frac{2}{3} \hat{s}_0^2$	-0.1553
$g_{LR}^{\nu d}$	$\frac{1}{3} \hat{s}_0^2$	0.0777
$g_{AV}^{ee}$	$\frac{1}{2} - 2 \hat{s}_0^2$	0.0225
$g_{AV}^{eu}$	$-\frac{1}{2} + \frac{4}{3} \hat{s}_0^2$	-0.1887
$g_{AV}^{ed}$	$\frac{1}{2} - \frac{2}{3} \hat{s}_0^2$	0.3419
$g_{VA}^{eu}$	$-\frac{1}{2} + 2 \hat{s}_0^2$	-0.0351
$g_{VA}^{ed}$	$\frac{1}{2} - 2 \hat{s}_0^2$	0.0247

The SM tree level expressions for the four-Fermi couplings are given in Table 10.3. Note that they differ from the respective products of the gauge couplings in Eq. (10.5) in the radiative corrections and in the presence of possible physics beyond the SM.

**10.3.1. Neutrino scattering:** For a general review on  $\nu$ -scattering we refer to Ref. 81 (nonstandard neutrino scattering interactions are surveyed in Ref. 82).

The cross-section in the laboratory system for  $\nu_\mu e \rightarrow \nu_\mu e$  or  $\bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$  elastic scattering [83] is (in this subsection we drop the redundant index  $L$  in the effective neutrino couplings)

$$\frac{d\sigma_{\nu, \bar{\nu}}}{dy} =$$

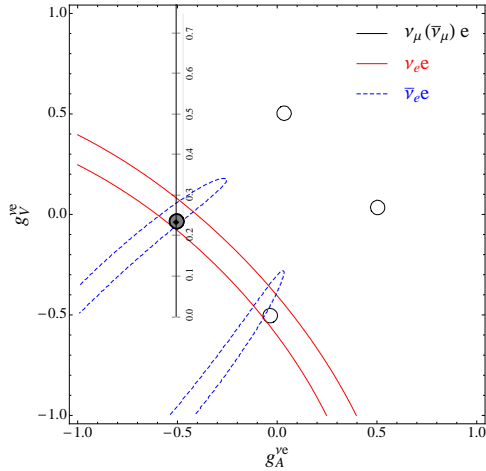
<sup>★</sup> We use here slightly different definitions (and to avoid confusion also a different notation) for the coefficients of these four-Fermi operators than we did in previous editions of this *Review*. The new couplings [80] are defined in the static limit,  $Q^2 \rightarrow 0$ , with specific radiative corrections included, while others (more experiment specific ones) are assumed to be removed by the experimentalist. They are convenient in that their determinations from very different types of processes can be straightforwardly combined.

$$\frac{G_F^2 m_e E_\nu}{2\pi} \left[ (g_V^{\nu e} \pm g_A^{\nu e})^2 + (g_V^{\nu e} \mp g_A^{\nu e})^2 (1-y)^2 - (g_V^{\nu e 2} - g_A^{\nu e 2}) \frac{y m_e}{E_\nu} \right], \quad (10.19)$$

where the upper (lower) sign refers to  $\nu_\mu(\bar{\nu}_\mu)$ , and  $y \equiv T_e/E_\nu$  (which runs from 0 to  $(1 + m_e/2E_\nu)^{-1}$ ) is the ratio of the kinetic energy of the recoil electron to the incident  $\nu$  or  $\bar{\nu}$  energy. For  $E_\nu \gg m_e$  this yields a total cross-section

$$\sigma = \frac{G_F^2 m_e E_\nu}{2\pi} \left[ (g_V^{\nu e} \pm g_A^{\nu e})^2 + \frac{1}{3} (g_V^{\nu e} \mp g_A^{\nu e})^2 \right]. \quad (10.20)$$

The most accurate measurements [83–88] of  $\sin^2 \theta_W$  from  $\nu$ -lepton scattering (see Sec. 10.6) are from the ratio  $R \equiv \sigma_{\nu\mu e}/\sigma_{\bar{\nu}\mu e}$ , in which many of the systematic uncertainties cancel. Radiative corrections (other than  $m_t$  effects) are small compared to the precision of present experiments and have negligible effect on the extracted  $\sin^2 \theta_W$ . The most precise experiment (CHARM II) [86] determined not only  $\sin^2 \theta_W$  but  $g_{V,A}^{\nu e}$  as well, which are shown in Fig. 10.1. The cross-sections for  $\nu_e e$  and  $\bar{\nu}_e e$  may be obtained from Eq. (10.19) by replacing  $g_{V,A}^{\nu e}$  by  $g_{V,A}^{\nu e} + 1$ , where the 1 is due to the charged-current contribution.



**Figure 10.1:** Allowed contours in  $g_A^{\nu e}$  vs.  $g_V^{\nu e}$  from neutrino-electron scattering and the SM prediction as a function of  $\hat{s}_Z^2$ . (The SM best fit value  $\hat{s}_Z^2 = 0.23129$  is also indicated.) The  $\nu_e e$  [87] and  $\bar{\nu}_e e$  [88] constraints are at 1  $\sigma$ , while each of the four equivalent  $\nu_\mu(\bar{\nu}_\mu)e$  [83–86] solutions ( $g_{V,A} \rightarrow -g_{V,A}$  and  $g_{V,A} \rightarrow g_{A,V}$ ) are at the 90% C.L. The global best fit region (shaded) almost exactly coincides with the corresponding  $\nu_\mu(\bar{\nu}_\mu)e$  region. The solution near  $g_A = 0, g_V = -0.5$  is eliminated by  $e^+e^- \rightarrow \ell^+\ell^-$  data under the weak additional assumption that the neutral current is dominated by the exchange of a single  $Z$  boson.

A precise determination of the on-shell  $s_W^2$ , which depends only very weakly on  $m_t$  and  $M_H$ , is obtained from deep inelastic scattering (DIS) of neutrinos from (approximately) isoscalar targets [89]. The ratio  $R_\nu \equiv \sigma_{\nu N}^{NC}/\sigma_{\nu N}^{CC}$  of neutral-to-charged-current cross-sections has been measured to 1% accuracy by CDHS [90] and CHARM [91] at CERN. CCFR [92] at Fermilab has obtained an even more precise result, so it is important to obtain theoretical expressions for  $R_\nu$  and  $R_{\bar{\nu}} \equiv \sigma_{\bar{\nu} N}^{NC}/\sigma_{\bar{\nu} N}^{CC}$  to comparable accuracy. Fortunately, many of the uncertainties from the strong interactions and neutrino spectra cancel in the ratio. A large theoretical uncertainty is associated with the  $c$ -threshold, which mainly affects  $\sigma^{CC}$ . Using the slow rescaling prescription [93] the central value of  $\sin^2 \theta_W$  from CCFR varies as  $0.0111(m_c/\text{GeV} - 1.31)$ , where  $m_c$  is the effective mass which is numerically close to the  $\overline{\text{MS}}$  mass  $\hat{m}_c(\hat{m}_c)$ , but their exact relation is unknown at higher orders. For  $m_c = 1.31 \pm 0.24$  GeV (determined from  $\nu$ -induced dimuon production [94]) this contributes  $\pm 0.003$  to the total uncertainty  $\Delta \sin^2 \theta_W \sim \pm 0.004$ . (The experimental

uncertainty is also  $\pm 0.003$ .) This uncertainty largely cancels, however, in the Paschos-Wolfenstein ratio [95],

$$R^- = \frac{\sigma_{\nu N}^{NC} - \sigma_{\bar{\nu} N}^{NC}}{\sigma_{\nu N}^{CC} - \sigma_{\bar{\nu} N}^{CC}}. \quad (10.21)$$

It was measured by Fermilab's NuTeV collaboration [96] for the first time, and required a high-intensity and high-energy anti-neutrino beam.

A simple zero<sup>th</sup>-order approximation is

$$R_\nu = g_L^2 + g_R^2 r, \quad R_{\bar{\nu}} = g_L^2 + \frac{g_R^2}{r}, \quad R^- = g_L^2 - g_R^2, \quad (10.22)$$

where

$$g_L^2 \equiv (g_{LL}^{\nu\mu u})^2 + (g_{LL}^{\nu\mu d})^2 \approx \frac{1}{2} - \sin^2 \theta_W + \frac{5}{9} \sin^4 \theta_W, \quad (10.23a)$$

$$g_R^2 \equiv (g_{LR}^{\nu\mu u})^2 + (g_{LR}^{\nu\mu d})^2 \approx \frac{5}{9} \sin^4 \theta_W, \quad (10.23b)$$

and  $r \equiv \sigma_{\bar{\nu} N}^{CC}/\sigma_{\nu N}^{CC}$  is the ratio of  $\bar{\nu}$  to  $\nu$  charged-current cross-sections, which can be measured directly. [In the simple parton model, ignoring hadron energy cuts,  $r \approx (\frac{1}{3} + \epsilon)/(\frac{1}{3} - \epsilon)$ , where  $\epsilon \sim 0.125$  is the ratio of the fraction of the nucleon's momentum carried by anti-quarks to that carried by quarks.] In practice, Eq. (10.22) must be corrected for quark mixing, quark sea effects,  $c$ -quark threshold effects, non-isoscalarity,  $W$ - $Z$  propagator differences, the finite muon mass, QED and EW radiative corrections. Details of the neutrino spectra, experimental cuts,  $x$  and  $Q^2$  dependence of structure functions, and longitudinal structure functions enter only at the level of these corrections and therefore lead to very small uncertainties. CCFR quotes  $s_W^2 = 0.2236 \pm 0.0041$  for  $(m_t, M_H) = (175, 150)$  GeV with very little sensitivity to  $(m_t, M_H)$ .

The NuTeV collaboration found  $s_W^2 = 0.2277 \pm 0.0016$  (for the same reference values), which was 3.0  $\sigma$  higher than the SM prediction [96]. The deviation was in  $g_L^2$  (initially 2.7  $\sigma$  low) while  $g_R^2$  was consistent with the SM. Since then a number of experimental and theoretical developments changed the interpretation of the measured cross-section ratios, affecting the extracted  $g_{L,R}^2$  (and thus  $s_W^2$ ) including their uncertainties and correlation. In the following paragraph we give a semi-quantitative and preliminary discussion of these effects, but we stress that the precise impact of them needs to be evaluated carefully by the collaboration with a new and self-consistent set of PDFs, including new radiative corrections, while simultaneously allowing isospin breaking and asymmetric strange seas. Until the time that such an effort is completed we do not include the  $\nu$ DIS constraints in our default set of fits.

(i) In the original analysis NuTeV worked with a symmetric strange quark sea but subsequently measured [97] the difference between the strange and antistrange momentum distributions,  $S^- \equiv \int_0^1 dx x[s(x) - \bar{s}(x)] = 0.00196 \pm 0.00143$ , from dimuon events utilizing the first complete next-to-leading order QCD description [98] and parton distribution functions (PDFs) according to Ref. 99. (ii) The measured branching ratio for  $K_{e3}$  decays enters crucially in the determination of the  $\nu_e(\bar{\nu}_e)$  contamination of the  $\nu_\mu(\bar{\nu}_\mu)$  beam. This branching ratio has moved from  $4.82 \pm 0.06\%$  at the time of the original publication [96] to the current value of  $5.07 \pm 0.04\%$ , *i.e.* a change by more than 4  $\sigma$ . This moves  $s_W^2$  about one standard deviation further away from the SM prediction while reducing the  $\nu_e(\bar{\nu}_e)$  uncertainty. (iii) PDFs seem to violate isospin symmetry at levels much stronger than generally expected [100]. A minimum  $\chi^2$  set of PDFs [101] allowing charge symmetry violation for both valence quarks [ $d_V^p(x) \neq u_V^n(x)$ ] and sea quarks [ $\bar{d}^p(x) \neq \bar{u}^n(x)$ ] shows a reduction in the NuTeV discrepancy by about 1  $\sigma$ . But isospin symmetry violating PDFs are currently not well constrained phenomenologically and within uncertainties the NuTeV anomaly could be accounted for in full or conversely made larger [101]. Still, the leading contribution from quark mass differences turns out to be largely model-independent [102] (at least in sign) and a shift,  $\delta s_W^2 = -0.0015 \pm 0.0003$  [103], has been estimated. (iv) QED splitting

effects also violate isospin symmetry with an effect on  $s_W^2$  whose sign (reducing the discrepancy) is model-independent. The corresponding shift of  $\delta s_W^2 = -0.0011$  has been calculated in Ref. 104 but has a large uncertainty. (v) Nuclear shadowing effects [105] are likely to affect the interpretation of the NuTeV result at some level, but the NuTeV collaboration argues that their data are dominated by values of  $Q^2$  at which nuclear shadowing is expected to be relatively small. However, another nuclear effect, known as the isovector EMC effect [106], is much larger (because it affects all neutrons in the nucleus, not just the excess ones) and model-independently works to reduce the discrepancy. It is estimated to lead to a shift of  $\delta s_W^2 = -0.0019 \pm 0.0006$  [103]. It would be important to verify and quantify this kind of effect experimentally, *e.g.*, in polarized electron scattering. (vi) The extracted  $s_W^2$  may also shift at the level of the quoted uncertainty when analyzed using the most recent QED and EW radiative corrections [107,108], as well as QCD corrections to the structure functions [109]. However, these are scheme-dependent and in order to judge whether they are significant they need to be adapted to the experimental conditions and kinematics of NuTeV, and have to be obtained in terms of observable variables and for the differential cross-sections. In addition, there is the danger of double counting some of the QED splitting effects. (vii) New physics could also affect  $g_{L,R}^2$  [110] but it is difficult to convincingly explain the entire effect that way.

**10.3.2. Parity violation:** For a review on weak polarized electron scattering we refer to Ref. 111. The SLAC polarized electron-deuteron DIS (eDIS) experiment [112] measured the right-left asymmetry,

$$A = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L}, \quad (10.24)$$

where  $\sigma_{R,L}$  is the cross-section for the deep-inelastic scattering of a right- or left-handed electron:  $e_{R,L}N \rightarrow eX$ . In the quark parton model,

$$\frac{A}{Q^2} = a_1 + a_2 \frac{1 - (1-y)^2}{1 + (1-y)^2}, \quad (10.25)$$

where  $Q^2 > 0$  is the momentum transfer and  $y$  is the fractional energy transfer from the electron to the hadrons. For the deuteron or other isoscalar targets, one has, neglecting the  $s$ -quark and anti-quarks,

$$a_1 = \frac{3G_F}{5\sqrt{2}\pi\alpha} \left( g_{AV}^{eu} - \frac{1}{2}g_{AV}^{ed} \right) \approx \frac{3G_F}{5\sqrt{2}\pi\alpha} \left( -\frac{3}{4} + \frac{5}{3}\hat{s}_0^2 \right) \quad (10.26a)$$

$$a_2 = \frac{3G_F}{5\sqrt{2}\pi\alpha} \left( g_{VA}^{eu} - \frac{1}{2}g_{VA}^{ed} \right) \approx \frac{9G_F}{5\sqrt{2}\pi\alpha} \left( \hat{s}_0^2 - \frac{1}{4} \right). \quad (10.26b)$$

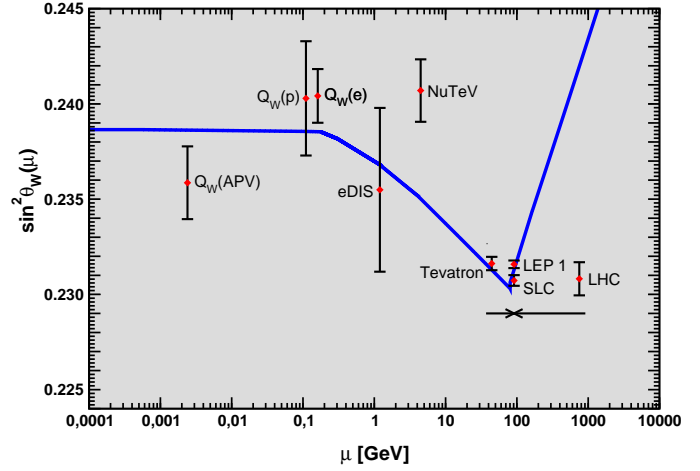
The Jefferson Lab Hall A Collaboration [113] improved on the SLAC result by determining  $A$  at  $Q^2 = 1.085$  GeV and 1.901 GeV, and determined the weak mixing angle to 2% precision. In another polarized-electron scattering experiment on deuterons, but in the quasi-elastic kinematic regime, the SAMPLE experiment [114] at MIT-Bates extracted the combination  $g_{VA}^{eu} - g_{VA}^{ed}$  at  $Q^2$  values of 0.1 GeV<sup>2</sup> and 0.038 GeV<sup>2</sup>. What was actually determined were nucleon form factors from which the quoted results were obtained by the removal of a multi-quark radiative correction [115]. Other linear combinations of the effective couplings have been determined in polarized-lepton scattering at CERN in  $\mu$ -C DIS, at Mainz in  $e$ -Be (quasi-elastic), and at Bates in  $e$ -C (elastic). See the review articles in Refs. 116 and 117 for more details. Recent polarized electron scattering experiments, *i.e.*, SAMPLE, the PVA4 experiment at Mainz, and the HAPPEX and GØ experiments at Jefferson Lab, have focussed on the strange quark content of the nucleon. These are reviewed in Refs. 118 and 119.

The parity violating asymmetry,  $A_{PV}$ , in fixed target polarized Møller scattering,  $e^-e^- \rightarrow e^-e^-$ , is defined as in Eq. (10.24) and reads [120],

$$\frac{A_{PV}}{Q^2} = -2g_{AV}^{ee} \frac{G_F}{\sqrt{2}\pi\alpha} \frac{1-y}{1+y^4 + (1-y)^4}, \quad (10.27)$$

where  $y$  is again the energy transfer. It has been measured at low  $Q^2 = 0.026$  GeV<sup>2</sup> in the SLAC E158 experiment [121], with

the result  $A_{PV} = (-1.31 \pm 0.14_{\text{stat.}} \pm 0.10_{\text{syst.}}) \times 10^{-7}$ . Expressed in terms of the weak mixing angle in the  $\overline{\text{MS}}$  scheme, this yields  $\hat{s}^2(Q^2) = 0.2403 \pm 0.0013$ , and established the scale dependence of the weak mixing angle (see  $Q_W(e)$  in Fig. 10.2) at the level of 6.4  $\sigma$ . One can also extract the model-independent effective coupling,  $g_{AV}^e = 0.0190 \pm 0.0027$  [80] (the implications are discussed in Ref. 123).



**Figure 10.2:** Scale dependence of the weak mixing angle defined in the  $\overline{\text{MS}}$  scheme [122] (for the scale dependence of the weak mixing angle defined in a mass-dependent renormalization scheme, see Ref. 123). The minimum of the curve corresponds to  $\mu = M_W$ , below which we switch to an effective theory with the  $W^\pm$  bosons integrated out, and where the  $\beta$ -function for the weak mixing angle changes sign. At the location of the  $W$  boson mass and each fermion mass there are also discontinuities arising from scheme dependent matching terms which are necessary to ensure that the various effective field theories within a given loop order describe the same physics. However, in the  $\overline{\text{MS}}$  scheme these are very small numerically and barely visible in the figure provided one decouples quarks at  $\mu = \hat{m}_q(\hat{m}_q)$ . The width of the curve reflects the theory uncertainty from strong interaction effects which at low energies is at the level of  $\pm 7 \times 10^{-5}$  [122]. Following the estimate [124] of the typical momentum transfer for parity violation experiments in Cs, the location of the APV data point is given by  $\mu = 2.4$  MeV. For NuTeV we display the updated value from Ref. 125 and chose  $\mu = \sqrt{20}$  GeV which is about half-way between the averages of  $\sqrt{Q^2}$  for  $\nu$  and  $\bar{\nu}$  interactions at NuTeV. The Tevatron and LHC measurements are strongly dominated by invariant masses of the final state dilepton pair of  $\mathcal{O}(M_Z)$  and can thus be considered as additional  $Z$  pole data points. For clarity we displayed the Tevatron and LHC points horizontally to the left and to the right, respectively.

In a similar experiment and at about the same  $Q^2 = 0.025$  GeV<sup>2</sup>, Qweak at Jefferson Lab [126] will be able to measure the weak charge of the proton (which is proportional to  $2g_{AV}^{eu} + g_{AV}^{ed}$ ) and  $\sin^2 \theta_W$  in polarized  $ep$  scattering with relative precisions of 4% and 0.3%, respectively. The result based on the collaborations commissioning run [127] and about 4% of the data corresponds to the constraint  $2g_{AV}^{eu} + g_{AV}^{ed} = 0.064 \pm 0.012$ .

There are precise experiments measuring atomic parity violation (APV) [128] in cesium [129,130] (at the 0.4% level [129]), thallium [131], lead [132], and bismuth [133]. The EW physics is contained in the nuclear weak charges  $Q_W^{Z,N}$ , where  $Z$  and  $N$  are the numbers of protons and neutrons in the nucleus. In terms of the nucleon vector couplings,

$$g_{AV}^{ep} \equiv 2g_{AV}^{eu} + g_{AV}^{ed} \approx -\frac{1}{2} + 2\hat{s}_0^2, \quad (10.28)$$

$$g_{AV}^{en} \equiv g_{AV}^{eu} + 2g_{AV}^{ed} \approx \frac{1}{2}, \quad (10.29)$$

one has,

$$Q_W^{Z,N} \equiv -2 [Z(g_{AV}^{ep} + 0.00005) + N(g_{AV}^{en} + 0.00006)] \left(1 - \frac{\alpha}{2\pi}\right), \quad (10.30)$$

where the numerically small adjustments are discussed in Ref. 80 and include the result of the  $\gamma Z$ -box correction from Ref. 134. *E.g.*,  $Q_W(^{133}\text{Cs})$  is extracted by measuring experimentally the ratio of the parity violating amplitude,  $E_{\text{PNC}}$ , to the Stark vector transition polarizability,  $\beta$ , and by calculating theoretically  $E_{\text{PNC}}$  in terms of  $Q_W$ . One can then write,

$$Q_W = N \left( \frac{\text{Im } E_{\text{PNC}}}{\beta} \right)_{\text{exp.}} \left( \frac{|e| a_B}{\text{Im } E_{\text{PNC}}} \frac{Q_W}{N} \right)_{\text{th.}} \left( \frac{\beta}{a_B^3} \right)_{\text{exp.+th.}} \left( \frac{a_B^2}{|e|} \right),$$

where  $a_B$  is the Bohr radius. The uncertainties associated with atomic wave functions are quite small for cesium [135]. The semi-empirical value of  $\beta$  used in early analyses added another source of theoretical uncertainty [136]. However, the ratio of the off-diagonal hyperfine amplitude to the polarizability was subsequently measured directly by the Boulder group [137]. Combined with the precisely known hyperfine amplitude [138] one finds  $\beta = (26.991 \pm 0.046) a_B^3$ , in excellent agreement with the earlier results, reducing the overall theory uncertainty (while slightly increasing the experimental error). Utilizing the state-of-the-art many-body calculation in Ref. 139 yields  $\text{Im } E_{\text{PNC}} = (0.8906 \pm 0.0026) \times 10^{-11} |e| a_B Q_W / N$ , while the two measurements [129,130] combine to give  $\text{Im } E_{\text{PNC}} / \beta = -1.5924 \pm 0.0055$  mV/cm, and we would obtain  $Q_W(^{133}\text{Cs}) = -73.20 \pm 0.35$ , or equivalently  $55g_{AV}^{ep} + 78g_{AV}^{en} = 36.64 \pm 0.18$  which is in excellent agreement with the SM prediction of 36.66. However, a very recent atomic structure calculation [140] found significant corrections to two non-dominating terms, changing the result to  $\text{Im } E_{\text{PNC}} = (0.8977 \pm 0.0040) \times 10^{-11} |e| a_B Q_W / N$ , and yielding the constraint,  $55g_{AV}^{ep} + 78g_{AV}^{en} = 36.35 \pm 0.21$  [ $Q_W(^{133}\text{Cs}) = -72.62 \pm 0.43$ ], *i.e.* a 1.5  $\sigma$  SM deviation. Thus, the various theoretical efforts in [139–141] together with an update of the SM calculation [142] reduced an earlier 2.3  $\sigma$  discrepancy from the SM (see the year 2000 edition of this *Review*), but there still appears to remain a small deviation. The theoretical uncertainties are 3% for thallium [143] but larger for the other atoms. The Boulder experiment in cesium also observed the parity-violating weak corrections to the nuclear electromagnetic vertex (the anapole moment [144]).

In the future it could be possible to further reduce the theoretical wave function uncertainties by taking the ratios of parity violation in different isotopes [128,145]. There would still be some residual uncertainties from differences in the neutron charge radii, however [146]. Experiments in hydrogen and deuterium are another possibility for reducing the atomic theory uncertainties [147], while measurements of single trapped radium ions are promising [148] because of the much larger parity violating effect.

#### 10.4. Physics of the massive electroweak bosons

If the CM energy  $\sqrt{s}$  is large compared to the fermion mass  $m_f$ , the unpolarized Born cross-section for  $e^+e^- \rightarrow f\bar{f}$  can be written as

$$\frac{d\sigma}{d\cos\theta} = \frac{\pi\alpha^2(s)}{2s} [F_1(1 + \cos^2\theta) + 2F_2\cos\theta] + B, \quad (10.31a)$$

where

$$F_1 = Q_e^2 Q_f^2 - 2\chi Q_e Q_f \bar{g}_V^e \bar{g}_V^f \cos\delta_R + \chi^2 (\bar{g}_V^e{}^2 + \bar{g}_A^e{}^2) (\bar{g}_V^f{}^2 + \bar{g}_A^f{}^2) \quad (10.31b)$$

$$F_2 = -2\chi Q_e Q_f \bar{g}_A^e \bar{g}_A^f \cos\delta_R + 4\chi^2 \bar{g}_V^e \bar{g}_A^e \bar{g}_V^f \bar{g}_A^f \quad (10.31c)$$

$$\tan\delta_R = \frac{\bar{M}_Z \bar{\Gamma}_Z}{\bar{M}_Z^2 - s}, \quad \chi = \frac{G_F}{2\sqrt{2}\pi\alpha(s)} \frac{s\bar{M}_Z^2}{[(\bar{M}_Z^2 - s)^2 + \bar{M}_Z^2 \bar{\Gamma}_Z^2]^{1/2}}, \quad (10.32)$$

and  $B$  accounts for box graphs involving virtual  $Z$  and  $W$  bosons, and  $\bar{g}_{V,A}^f$  are defined in Eq. (10.33) below.  $\bar{M}_Z$  and  $\bar{\Gamma}_Z$  correspond to mass and width definitions based on a Breit-Wigner shape with

an energy-independent width (see the Section on “The  $Z$  Boson” in the Gauge and Higgs Boson Particle Listings of this *Review*). The differential cross-section receives important corrections from QED effects in the initial and final state, and interference between the two (see *e.g.* Ref. 149). For  $q\bar{q}$  production, there are additional final-state QCD corrections, which are relatively large. Note also that the equations above are written in the CM frame of the incident  $e^+e^-$  system, which may be boosted due to the initial-state QED radiation.

Some of the leading virtual EW corrections are captured by the running QED coupling  $\alpha(s)$  and the Fermi constant  $G_F$ . The remaining corrections to the  $Zf\bar{f}$  interaction are absorbed by replacing the tree-level couplings in Eq. (10.5) with the  $s$ -dependent *effective couplings* [150],

$$\bar{g}_V^f = \sqrt{\rho_f} (t_{3L}^{(f)} - 2Q_f \kappa_f \sin^2\theta_W), \quad \bar{g}_A^f = \sqrt{\rho_f} t_{3L}^{(f)}. \quad (10.33)$$

In these equations, the effective couplings are to be taken at the scale  $\sqrt{s}$ , but for notational simplicity we do not show this explicitly. At tree-level  $\rho_f = \kappa_f = 1$ , but inclusion of EW radiative corrections leads to non-zero  $\rho_f - 1$  and  $\kappa_f - 1$ , which depend on the fermion  $f$  and on the renormalization scheme. In the on-shell scheme, the quadratic  $m_t$  dependence is given by  $\rho_f \sim 1 + \rho_t$ ,  $\kappa_f \sim 1 + \rho_t / \tan^2\theta_W$ , while in  $\overline{\text{MS}}$ ,  $\bar{\rho}_f \sim \bar{\kappa}_f \sim 1$ , for  $f \neq b$  ( $\bar{\rho}_b \sim 1 - \frac{4}{3}\rho_t$ ,  $\bar{\kappa}_b \sim 1 + \frac{2}{3}\rho_t$ ). In the  $\overline{\text{MS}}$  scheme the normalization is changed according to  $G_F M_Z^2 / 2\sqrt{2}\pi \rightarrow \hat{\alpha} / 4\hat{s}_Z^2 \hat{c}_Z^2$  in Eq. (10.32).

For the high-precision  $Z$ -pole observables discussed below, additional bosonic and fermionic loops, vertex corrections, and higher order contributions, *etc.*, must be included [67,68,151,152,153]. For example, in the  $\overline{\text{MS}}$  scheme one has  $\hat{\rho}_\ell = 0.9980$ ,  $\hat{\kappa}_\ell = 1.0010$ ,  $\bar{\rho}_b = 0.9868$ , and  $\bar{\kappa}_b = 1.0065$ .

To connect to measured quantities, it is convenient to define an effective angle  $\bar{s}_Z^2 \equiv \sin^2\bar{\theta}_{Wf} \equiv \hat{\kappa}_f \hat{s}_Z^2 = \kappa_f s_W^2$ , in terms of which  $\bar{g}_V^f$  and  $\bar{g}_A^f$  are given by  $\sqrt{\rho_f}$  times their tree-level formulae. One finds that the  $\hat{\kappa}_f$  ( $f \neq b$ ) are almost independent of  $(m_t, M_H)$ , and thus one can write

$$\bar{s}_\ell^2 = \bar{s}_Z^2 + 0.00023, \quad (10.34)$$

while the  $\kappa$ 's for the other schemes are  $m_t$  dependent.

##### 10.4.1. $e^+e^-$ scattering below the $Z$ pole :

Experiments at PEP, PETRA and TRISTAN have measured the unpolarized forward-backward asymmetry,  $A_{FB}$ , and the total cross-section relative to pure QED,  $R$ , for  $e^+e^- \rightarrow \ell^+\ell^-$ ,  $\ell = \mu$  or  $\tau$  at CM energies  $\sqrt{s} < M_Z$ . They are defined as

$$A_{FB} \equiv \frac{\sigma_F - \sigma_B}{\sigma_F + \sigma_B}, \quad R = \frac{\sigma}{\mathcal{R}_{\text{ini}} 4\pi\alpha^2/3s}, \quad (10.35)$$

where  $\sigma_F$  ( $\sigma_B$ ) is the cross-section for  $\ell^-$  to travel forward (backward) with respect to the  $e^-$  direction. Neglecting box graph contribution, they are given by

$$A_{FB} = \frac{3F_2}{4F_1}, \quad R = F_1. \quad (10.36)$$

For the available data, it is sufficient to approximate the EW corrections through the leading running  $\alpha(s)$  and quadratic  $m_t$  contributions [154,155] as described above. Reviews and formulae for  $e^+e^- \rightarrow \text{hadrons}$  may be found in Ref. 156.

##### 10.4.2. $Z$ pole physics :

High-precision measurements of various  $Z$  pole ( $\sqrt{s} \approx M_Z$ ) observables have been performed at LEP 1 and SLC [10,157–162], as summarized in Table 10.5. These include the  $Z$  mass and total width,  $\Gamma_Z$ , and partial widths  $\Gamma(f\bar{f})$  for  $Z \rightarrow f\bar{f}$ , where  $f = e, \mu, \tau$ , light hadrons,  $b$ , or  $c$ . It is convenient to use the variables  $M_Z$ ,  $\Gamma_Z$ ,  $R_\ell \equiv \Gamma(\text{had})/\Gamma(\ell^+\ell^-)$  ( $\ell = e, \mu, \tau$ ),  $\sigma_{\text{had}} \equiv 12\pi\Gamma(e^+e^-)\Gamma(\text{had})/M_Z^2\Gamma_Z^2$ ††,

†† Note that  $\sigma_{\text{had}}$  receives additional EW corrections that are not captured in the partial widths [163,153], but they only enter at two-loop order.



$R_b \equiv \Gamma(b\bar{b})/\Gamma(\text{had})$ , and  $R_c \equiv \Gamma(c\bar{c})/\Gamma(\text{had})$ , most of which are weakly correlated experimentally. ( $\Gamma(\text{had})$  is the partial width into hadrons.) The three values for  $R_\ell$  are consistent with lepton universality (although  $R_\tau$  is somewhat low compared to  $R_e$  and  $R_\mu$ ), but we use the general analysis in which the three observables are treated as independent. Similar remarks apply to  $A_{FB}^{0,\ell}$  defined through Eq. (10.39) with  $P_e = 0$  ( $A_{FB}^{0,\tau}$  is somewhat high).  $\mathcal{O}(\alpha^3)$  QED corrections introduce a large anti-correlation ( $-30\%$ ) between  $\Gamma_Z$  and  $\sigma_{\text{had}}$ . The anti-correlation between  $R_b$  and  $R_c$  is  $-18\%$  [10]. The  $R_\ell$  are insensitive to  $m_t$  except for the  $Z \rightarrow b\bar{b}$  vertex and final state corrections and the implicit dependence through  $\sin^2 \theta_W$ . Thus, they are especially useful for constraining  $\alpha_s$ . The invisible decay width [10],  $\Gamma(\text{inv}) = \Gamma_Z - 3\Gamma(\ell^+\ell^-) - \Gamma(\text{had}) = 499.0 \pm 1.5$  MeV, can be used to determine the number of neutrino flavors,  $N_\nu = \Gamma(\text{inv})/\Gamma^{\text{theory}}(\nu\bar{\nu})$ , much lighter than  $M_Z/2$ . In practice, we determine  $N_\nu$  by allowing it as an additional fit parameter and obtain,

$$N_\nu = 2.992 \pm 0.007. \quad (10.37)$$

Additional constraints follow from measurements of various  $Z$ -pole asymmetries. These include the forward-backward asymmetry  $A_{FB}$  and the polarization or left-right asymmetry,

$$A_{LR} \equiv \frac{\sigma_L - \sigma_R}{\sigma_L + \sigma_R}, \quad (10.38)$$

where  $\sigma_L(\sigma_R)$  is the cross-section for a left-(right-)handed incident electron.  $A_{LR}$  was measured precisely by the SLD collaboration at the SLC [159], and has the advantages of being very sensitive to  $\sin^2 \theta_W$  and that systematic uncertainties largely cancel. After removing initial state QED corrections and contributions from photon exchange,  $\gamma$ - $Z$  interference and EW boxes, see Eq. (10.31), one can use the effective tree-level expressions

$$A_{LR} = A_e P_e, \quad A_{FB} = \frac{3}{4} A_f \frac{A_e + P_e}{1 + P_e A_e}, \quad (10.39)$$

where

$$A_f \equiv \frac{2\bar{g}_V^f \bar{g}_A^f}{\bar{g}_V^f{}^2 + \bar{g}_A^f{}^2} = \frac{1 - 4|Q_f|\bar{s}_f^2}{1 - 4|Q_f|\bar{s}_f^2 + 8(|Q_f|\bar{s}_f^2)^2}. \quad (10.40)$$

$P_e$  is the initial  $e^-$  polarization, so that the second equality in Eq. (10.41) is reproduced for  $P_e = 1$ , and the  $Z$  pole forward-backward asymmetries at LEP 1 ( $P_e = 0$ ) are given by  $A_{FB}^{(0,f)} = \frac{3}{4} A_e A_f$  where  $f = e, \mu, \tau, b, c, s$  [10], and  $q$ , and where  $A_{FB}^{(0,q)}$  refers to the hadronic charge asymmetry. Corrections for  $t$ -channel exchange and  $s/t$ -channel interference cause  $A_{FB}^{(0,e)}$  to be strongly anti-correlated with  $R_e$  ( $-37\%$ ). The correlation between  $A_{FB}^{(0,b)}$  and  $A_{FB}^{(0,c)}$  amounts to 15%.

In addition, SLD extracted the final-state couplings  $A_b, A_c$  [10],  $A_s$  [160],  $A_\tau$ , and  $A_\mu$  [161], from left-right forward-backward asymmetries, using

$$A_{LR}^{FB}(f) = \frac{\sigma_{LF}^f - \sigma_{LB}^f - \sigma_{RF}^f + \sigma_{RB}^f}{\sigma_{LF}^f + \sigma_{LB}^f + \sigma_{RF}^f + \sigma_{RB}^f} = \frac{3}{4} A_f, \quad (10.41)$$

where, for example,  $\sigma_{LF}^f$  is the cross-section for a left-handed incident electron to produce a fermion  $f$  traveling in the forward hemisphere. Similarly,  $A_\tau$  and  $A_e$  were measured at LEP 1 [10] through the  $\tau$  polarization,  $\mathcal{P}_\tau$ , as a function of the scattering angle  $\theta$ , which can be written as

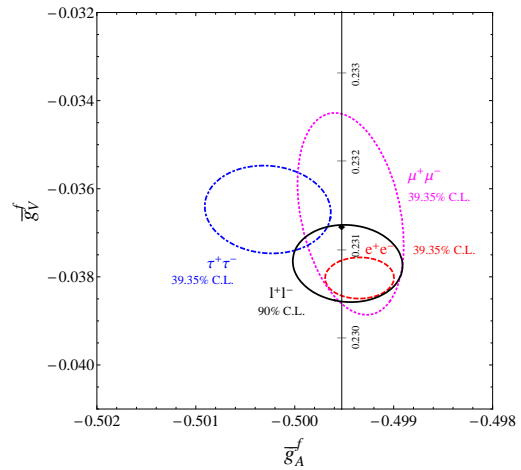
$$\mathcal{P}_\tau = -\frac{A_\tau(1 + \cos^2 \theta) + 2A_e \cos \theta}{(1 + \cos^2 \theta) + 2A_\tau A_e \cos \theta} \quad (10.42)$$

The average polarization,  $\langle \mathcal{P}_\tau \rangle$ , obtained by integrating over  $\cos \theta$  in the numerator and denominator of Eq. (10.42), yields  $\langle \mathcal{P}_\tau \rangle = -A_\tau$ , while  $A_e$  can be extracted from the angular distribution of  $\mathcal{P}_\tau$ .

The initial state coupling,  $A_e$ , was also determined through the left-right charge asymmetry [162] and in polarized Bhabba scattering [161]

at SLC. Because  $\bar{g}_V^\ell$  is very small, not only  $A_{LR}^0 = A_e$ ,  $A_{FB}^{(0,\ell)}$ , and  $\mathcal{P}_\tau$ , but also  $A_{FB}^{(0,b)}$ ,  $A_{FB}^{(0,c)}$ ,  $A_{FB}^{(0,s)}$ , and the hadronic asymmetries are mainly sensitive to  $\bar{s}_\ell^2$ .

As mentioned in Sec. 10.2, radiative corrections to  $\bar{s}_\ell^2$  have been computed with full two-loop and partial higher-order corrections. Moreover, fermionic two-loop EW corrections to  $\bar{s}_q^2$  ( $q = b, c, s$ ) have been obtained [79,152], but the purely bosonic contributions of this order are still missing. Similarly, for the partial widths,  $\Gamma(f\bar{f})$ , and the hadronic peak cross-section,  $\sigma_{\text{had}}$ , the fermionic two-loop EW corrections are known [153]. Non-factorizable  $\mathcal{O}(\alpha\alpha_s)$  corrections to the  $Z \rightarrow q\bar{q}$  vertex are also available [151]. They add coherently, resulting in a sizable effect and shift  $\alpha_s(M_Z)$  when extracted from  $Z$  lineshape observables by  $\approx +0.0007$ . As an example of the precision of the  $Z$ -pole observables, the values of  $\bar{g}_A^f$  and  $\bar{g}_V^f$ ,  $f = e, \mu, \tau, \ell$ , extracted from the LEP and SLC lineshape and asymmetry data, are shown in Fig. 10.3, which should be compared with Fig. 10.1. (The two sets of parameters coincide in the SM at tree-level.)



**Figure 10.3:**  $1\sigma$  (39.35% C.L.) contours for the  $Z$ -pole observables  $\bar{g}_A^f$  and  $\bar{g}_V^f$ ,  $f = e, \mu, \tau$  obtained at LEP and SLC [10], compared to the SM expectation as a function of  $\hat{s}_Z^2$ . (The SM best fit value  $\hat{s}_Z^2 = 0.23129$  is also indicated.) Also shown is the 90% CL allowed region in  $\bar{g}_{A,V}^\ell$  obtained assuming lepton universality.

As for hadron colliders, the forward-backward asymmetry,  $A_{FB}$ , for  $e^+e^-$  and  $\mu^+\mu^-$  final states (with invariant masses restricted to or dominated by values around  $M_Z$ ) in  $p\bar{p}$  collisions has been measured by the DØ [164] (only  $e^+e^-$ ) and CDF [165,166] collaborations, and the values  $\bar{s}_\ell^2 = 0.23146 \pm 0.00047$  and  $\bar{s}_\ell^2 = 0.23222 \pm 0.00046$ , were extracted, respectively. Assuming that the smaller systematic uncertainty ( $\pm 0.00018$  from CDF [166]) is common to both experiments, these measurements combine to

$$\bar{s}_\ell^2 = 0.23185 \pm 0.00035 \text{ (Tevatron)} \quad (10.43)$$

By varying the invariant mass and the scattering angle (and assuming the electron couplings), information on the effective  $Z$  couplings to light quarks,  $\bar{g}_{V,A}^{u,d}$ , could also be obtained [167,168], but with large uncertainties and mutual correlations and not independently of  $\bar{s}_\ell^2$  above. Similar analyses have also been reported by the H1 and ZEUS collaborations at HERA [169] and by the LEP collaborations [10]. This kind of measurement is harder in the  $pp$  environment due to the difficulty to assign the initial quark and antiquark in the underlying Drell-Yan process to the protons. Nevertheless, measurements of  $A_{FB}$  have been reported by the ATLAS [170], CMS [171] and LHCb [172] collaborations (the latter two only for the  $\mu^+\mu^-$  final state), which obtained  $\bar{s}_\ell^2 = 0.2308 \pm 0.0012$ ,  $\bar{s}_\ell^2 = 0.2287 \pm 0.0032$  and  $\bar{s}_\ell^2 = 0.23142 \pm 0.00106$ , respectively. Assuming that the smallest

theoretical uncertainty ( $\pm 0.00056$  from LHCb [172]) is fully correlated among all three experiments, these measurements combine to

$$\bar{s}_\ell^2 = 0.23105 \pm 0.00087 \text{ (LHC)} \quad (10.44)$$

#### 10.4.3. LEP 2 :

LEP 2 [173,174] ran at several energies above the  $Z$  pole up to  $\sim 209$  GeV. Measurements were made of a number of observables, including the cross-sections for  $e^+e^- \rightarrow f\bar{f}$  for  $f = q, \mu, \tau$ ; the differential cross-sections for  $f = e, \mu, \tau$ ;  $R_q$  for  $q = b, c$ ;  $A_{FB}(f)$  for  $f = \mu, \tau, b, c$ ;  $W$  branching ratios; and  $\gamma\gamma$ ,  $WW$ ,  $WW\gamma$ ,  $ZZ$ , single  $W$ , and single  $Z$  cross-sections. They are in good agreement with the SM predictions, with the exceptions of  $R_b$  ( $2.1 \sigma$  low),  $A_{FB}(b)$  ( $1.6 \sigma$  low), and the  $W \rightarrow \tau\nu_\tau$  branching fraction ( $2.6 \sigma$  high).

The  $Z$  boson properties are extracted assuming the SM expressions for the  $\gamma$ - $Z$  interference terms. These have also been tested experimentally by performing more general fits [173,175] to the LEP 1 and LEP 2 data. Assuming family universality this approach introduces three additional parameters relative to the standard fit [10], describing the  $\gamma$ - $Z$  interference contribution to the total hadronic and leptonic cross-sections,  $j_{\text{had}}^{\text{tot}}$  and  $j_\ell^{\text{tot}}$ , and to the leptonic forward-backward asymmetry,  $j_\ell^{\text{fb}}$ . *E.g.*,

$$j_{\text{had}}^{\text{tot}} \sim g_V^\ell g_V^{\text{had}} = 0.277 \pm 0.065, \quad (10.45)$$

which is in agreement with the SM expectation [10] of  $0.21 \pm 0.01$ . These are valuable tests of the SM; but it should be cautioned that new physics is not expected to be described by this set of parameters, since (i) they do not account for extra interactions beyond the standard weak neutral current, and (ii) the photonic amplitude remains fixed to its SM value.

Strong constraints on anomalous triple and quartic gauge couplings have been obtained at LEP 2 and the Tevatron as described in the Gauge & Higgs Bosons Particle Listings.

#### 10.4.4. $W$ and $Z$ decays :

The partial decay widths for gauge bosons to decay into massless fermions  $f_1\bar{f}_2$  (the numerical values include the small EW radiative corrections and final state mass effects) are given by

$$\Gamma(W^+ \rightarrow e^+\nu_e) = \frac{G_F M_W^3}{6\sqrt{2}\pi} \approx 226.27 \pm 0.05 \text{ MeV}, \quad (10.46a)$$

$$\begin{aligned} \Gamma(W^+ \rightarrow u_i\bar{d}_j) &= \frac{\mathcal{R}_V^q G_F M_W^3}{6\sqrt{2}\pi} |V_{ij}|^2 \\ &\approx 705.1 \pm 0.3 \text{ MeV} |V_{ij}|^2, \end{aligned} \quad (10.46b)$$

$$\begin{aligned} \Gamma(Z \rightarrow f\bar{f}) &= \frac{G_F M_Z^3}{6\sqrt{2}\pi} [\mathcal{R}_V^f g_V^{f2} + \mathcal{R}_A^f g_A^{f2}] \\ &\approx \begin{cases} 167.17 \pm 0.02 \text{ MeV} (\nu\bar{\nu}), \\ 83.97 \pm 0.01 \text{ MeV} (e^+e^-), \\ 299.91 \pm 0.19 \text{ MeV} (u\bar{u}), \\ 382.80 \pm 0.14 \text{ MeV} (d\bar{d}), \\ 375.69 \mp 0.17 \text{ MeV} (b\bar{b}). \end{cases} \end{aligned} \quad (10.46c)$$

Final-state QED and QCD corrections to the vector and axial-vector form factors are given by

$$\mathcal{R}_{V,A}^f = N_C [1 + \frac{3}{4} (Q_f^2 \frac{\alpha(s)}{\pi} + \frac{N_C^2 - 1}{2N_C} \frac{\alpha_s(s)}{\pi}) + \dots], \quad (10.47)$$

where  $N_C = 3$  (1) is the color factor for quarks (leptons) and the dots indicate finite fermion mass effects proportional to  $m_f^2/s$  which are different for  $\mathcal{R}_V^f$  and  $\mathcal{R}_A^f$ , as well as higher-order QCD corrections, which are known to  $\mathcal{O}(\alpha_s^4)$  [176–178]. These include singlet contributions starting from two-loop order which are large, strongly top quark mass dependent, family universal, and flavor non-universal [179]. Also the  $\mathcal{O}(\alpha^2)$  self-energy corrections from Ref. 180 are taken into account.

For the  $W$  decay into quarks, Eq. (10.46b), only the universal massless part (non-singlet and  $m_q = 0$ ) of the final-state QCD radiator function in  $\mathcal{R}_V$  from Eq. (10.47) is used, and the QED corrections are modified. Expressing the widths in terms of  $G_F M_{W,Z}^3$  incorporates the largest radiative corrections from the running QED coupling [56,181]. EW corrections to the  $Z$  widths are then taken into account through the effective couplings  $\bar{g}_{V,A}^{i2}$ . Hence, in the on-shell scheme the  $Z$  widths are proportional to  $\rho_i \sim 1 + \rho_t$ . There is additional (negative) quadratic  $m_t$  dependence in the  $Z \rightarrow b\bar{b}$  vertex corrections [182] which causes  $\Gamma(b\bar{b})$  to decrease with  $m_t$ . The dominant effect is to multiply  $\Gamma(b\bar{b})$  by the vertex correction  $1 + \delta\rho_{b\bar{b}}$ , where  $\delta\rho_{b\bar{b}} \sim 10^{-2}(-\frac{1}{2}m_t^2/M_Z^2 + \frac{1}{5})$ . In practice, the corrections are included in  $\rho_b$  and  $\kappa_b$ , as discussed in Sec. 10.4.

For three fermion families the total widths are predicted to be

$$\Gamma_Z \approx 2.4943 \pm 0.0008 \text{ GeV}, \quad \Gamma_W \approx 2.0888 \pm 0.0007 \text{ GeV}. \quad (10.48)$$

The uncertainties in these predictions are almost entirely induced from the fit error in  $\alpha_s(M_Z) = 0.1182 \pm 0.0016$ . These predictions are to be compared with the experimental results,  $\Gamma_Z = 2.4952 \pm 0.0023 \text{ GeV}$  [10] and  $\Gamma_W = 2.085 \pm 0.042 \text{ GeV}$  (see the Gauge & Higgs Boson Particle Listings for more details).

#### 10.4.5. $H$ decays :

The ATLAS and CMS collaborations at LHC observed a Higgs boson [183] with properties appearing well consistent with the SM Higgs (see the note on “The Higgs Boson  $H^0$ ” in the Gauge & Higgs Boson Particle Listings). A recent combination [184] of ATLAS and CMS results for the Higgs boson mass from kinematical reconstruction yields

$$M_H = 125.09 \pm 0.24 \text{ GeV}. \quad (10.49)$$

In analogy to the  $W$  and  $Z$  decays discussed in the previous subsection, we can include some of the Higgs decay properties into the global analysis of Sec. 10.6. However, the total Higgs decay width, which in the SM amounts to

$$\Gamma_H = 4.15 \pm 0.06 \text{ MeV}, \quad (10.50)$$

is too small to be resolved at the LHC. However, one can employ results of Higgs branching ratios into different final states. The most useful channels are Higgs decays into  $WW^*$  and  $ZZ^*$  (with at least one gauge boson off-shell), as well as  $\gamma\gamma$  and  $\tau\tau$ . We define

$$\rho_{XY} \equiv \ln \frac{\text{BR}_{H \rightarrow XX}}{\text{BR}_{H \rightarrow YY}}. \quad (10.51)$$

These quantities are constructed to have a SM expectation of zero, and their physical range is over all real numbers, which allows one to straightforwardly use Gaussian error propagation (in view of the fairly large errors). Moreover, possible effects of new physics on Higgs production rates would also cancel and one may focus on the decay side of the processes. From a combination of ATLAS and CMS results [184], we find

$$\rho_{\gamma W} = -0.03 \pm 0.20, \quad \rho_{\tau Z} = -0.27 \pm 0.31,$$

which we take to be uncorrelated as they involve distinct final states. We evaluate the decay rates with the package HDECAY [185].

### 10.5. Precision flavor physics

In addition to cross-sections, asymmetries, parity violation,  $W$  and  $Z$  decays, there is a large number of experiments and observables testing the flavor structure of the SM. These are addressed elsewhere in this *Review*, and are generally not included in this Section. However, we identify three precision observables with sensitivity to similar types of new physics as the other processes discussed here. The branching fraction of the flavor changing transition  $b \rightarrow s\gamma$  is of comparatively low precision, but since it is a loop-level process (in the SM) its sensitivity to new physics (and SM parameters, such as heavy quark masses) is enhanced. A discussion can be found in the 2010 edition of this *Review*. The  $\tau$ -lepton lifetime and leptonic branching

ratios are primarily sensitive to  $\alpha_s$  and not affected significantly by many types of new physics. However, having an independent and reliable low energy measurement of  $\alpha_s$  in a global analysis allows the comparison with the  $Z$  lineshape determination of  $\alpha_s$  which shifts easily in the presence of new physics contributions. By far the most precise observable discussed here is the anomalous magnetic moment of the muon (the electron magnetic moment is measured to even greater precision and can be used to determine  $\alpha$ , but its new physics sensitivity is suppressed by an additional factor of  $m_e^2/m_\mu^2$ , unless there is a new light degree of freedom such as a dark  $Z$  [186] boson). Its combined experimental and theoretical uncertainty is comparable to typical new physics contributions.

The extraction of  $\alpha_s$  from the  $\tau$  lifetime [187] is standing out from other determinations because of a variety of independent reasons: (i) the  $\tau$ -scale is low, so that upon extrapolation to the  $Z$  scale (where it can be compared to the theoretically clean  $Z$  lineshape determinations) the  $\alpha_s$  error shrinks by about an order of magnitude; (ii) yet, this scale is high enough that perturbation theory and the operator product expansion (OPE) can be applied; (iii) these observables are fully inclusive and thus free of fragmentation and hadronization effects that would have to be modeled or measured; (iv) duality violation (DV) effects are most problematic near the branch cut but there they are suppressed by a double zero at  $s = m_\tau^2$ ; (v) there are data [34,188] to constrain non-perturbative effects both within and breaking the OPE; (vi) a complete four-loop order QCD calculation is available [178]; (vii) large effects associated with the QCD  $\beta$ -function can be re-summed [189] in what has become known as contour improved perturbation theory (CIPT). However, while there is no doubt that CIPT shows faster convergence in the lower (calculable) orders, doubts have been cast on the method by the observation that at least in a specific model [190], which includes the exactly known coefficients and theoretical constraints on the large-order behavior, ordinary fixed order perturbation theory (FOPT) may nevertheless give a better approximation to the full result. We therefore use the expressions [45,177,178,191],

$$\tau_\tau = \hbar \frac{1 - \mathcal{B}_\tau^s}{\Gamma_\tau^e + \Gamma_\tau^\mu + \Gamma_\tau^{ud}} = 290.88 \pm 0.35 \text{ fs}, \quad (10.52)$$

$$\begin{aligned} \Gamma_\tau^{ud} = & \frac{G_F^2 m_\tau^5 |V_{ud}|^2}{64\pi^3} S(m_\tau, M_Z) \left( 1 + \frac{3}{5} \frac{m_\tau^2 - m_\mu^2}{M_W^2} \right) \times \\ & \left[ 1 + \frac{\alpha_s(m_\tau)}{\pi} + 5.202 \frac{\alpha_s^2}{\pi^2} + 26.37 \frac{\alpha_s^3}{\pi^3} + \right. \\ & \left. 127.1 \frac{\alpha_s^4}{\pi^4} + \frac{\hat{\alpha}}{\pi} \left( \frac{85}{24} - \frac{\pi^2}{2} \right) + \delta_{\text{NP}} \right], \quad (10.53) \end{aligned}$$

and  $\Gamma_\tau^e$  and  $\Gamma_\tau^\mu$  can be taken from Eq. (10.6) with obvious replacements. The relative fraction of decays with  $\Delta S = -1$ ,  $\mathcal{B}_\tau^s = 0.0287 \pm 0.0005$ , is based on experimental data since the value for the strange quark mass,  $\hat{m}_s(m_\tau)$ , is not well known and the QCD expansion proportional to  $\hat{m}_s^2$  converges poorly and cannot be trusted.  $S(m_\tau, M_Z) = 1.01907 \pm 0.0003$  is a logarithmically enhanced EW correction factor with higher orders re-summed [192].  $\delta_{\text{NP}}$  collects non-perturbative and quark-mass suppressed contributions, including the dimension four, six and eight terms in the OPE, as well as DV effects. We use the average,  $\delta_{\text{NP}} = 0.0114 \pm 0.0072$ , of the slightly conflicting results,  $\delta_{\text{NP}} = -0.004 \pm 0.012$  [193] and  $\delta_{\text{NP}} = 0.020 \pm 0.0009$  [194], based on OPAL [34] and ALEPH [188]  $\tau$  spectral functions, respectively. The dominant uncertainty arises from the truncation of the FOPT series and is conservatively taken as the  $\alpha_s^4$  term (this is re-calculated in each call of the fits, leading to an  $\alpha_s$ -dependent and thus asymmetric error) until a better understanding of the numerical differences between FOPT and CIPT has been gained. Our perturbative error covers almost the entire range from using CIPT to assuming that the nearly geometric series in Eq. (10.53) continues to higher orders. The experimental uncertainty in Eq. (10.52), is from the combination of the two leptonic branching ratios with the direct  $\tau_\tau$ . Included are also various smaller uncertainties ( $\pm 0.5$  fs) from other sources which are dominated by the evolution from the  $Z$  scale. In total we obtain a  $\sim 1.5\%$  determination of  $\alpha_s(M_Z) = 0.1174_{-0.0017}^{+0.0019}$ ,

which corresponds to  $\alpha_s(m_\tau) = 0.314_{-0.013}^{+0.016}$ , and updates the result of Refs. 45 and 195. For more details, see Refs. 193 and 194 where the  $\tau$  spectral functions themselves and an estimate of the unknown  $\alpha_s^5$  term are used as additional inputs.

The world average of the muon anomalous magnetic moment<sup>†</sup>,

$$a_\mu^{\text{exp}} = \frac{g_\mu - 2}{2} = (1165920.91 \pm 0.63) \times 10^{-9}, \quad (10.54)$$

is dominated by the final result of the E821 collaboration at BNL [196]. The QED contribution has been calculated to five loops [197] (fully analytic to three loops [198,199]). The estimated SM EW contribution [200–202],  $a_\mu^{\text{EW}} = (1.54 \pm 0.01) \times 10^{-9}$ , which includes two-loop [201] and leading three-loop [202] corrections, is at the level of twice the current uncertainty.

The limiting factor in the interpretation of the result are the uncertainties from the two- and three-loop hadronic contribution [203]. *E.g.*, Ref. 31 obtained the value  $a_\mu^{\text{had}} = (69.23 \pm 0.42) \times 10^{-9}$  which combines CMD-2 [38] and SND [39]  $e^+e^- \rightarrow \text{hadrons}$  cross-section data with radiative return results from BaBar [41] and KLOE [42]. The most recent analysis [20] includes  $\tau$  decay data corrected for isospin symmetry violation and  $\gamma$ - $\rho$  mixing effects and yields  $a_\mu^{\text{had}} = (68.72 \pm 0.35) \times 10^{-9}$ . The largest isospin symmetry violating effect is due to higher-order EW corrections [204] but introduces a negligible uncertainty [192]. The  $\gamma$ - $\rho$  mixing effect [205] resolves an earlier discrepancy between the spectral functions obtained from  $e^+e^-$  and  $\tau$  decay data. A recent lattice QCD calculation finds agreement with the results of the dispersive approach discussed above although within a much larger error [206].

An additional uncertainty is induced by the hadronic three-loop light-by-light scattering contribution. Several recent independent model calculations yield compatible results:  $a_\mu^{\text{LBS}}(\alpha^3) = (+1.36 \pm 0.25) \times 10^{-9}$  [207],  $a_\mu^{\text{LBS}}(\alpha^3) = +1.37_{-0.27}^{+0.15} \times 10^{-9}$  [208],  $a_\mu^{\text{LBS}}(\alpha^3) = (+1.16 \pm 0.40) \times 10^{-9}$  [209], and  $a_\mu^{\text{LBS}}(\alpha^3) = (+1.05 \pm 0.26) \times 10^{-9}$  [210]. The sign of this effect is opposite [211] to the one quoted in the 2002 edition of this *Review*, and its magnitude is larger than previous evaluations [211,212]. There is also an upper bound  $a_\mu^{\text{LBS}}(\alpha^3) < 1.59 \times 10^{-9}$  [208] but this requires an *ad hoc* assumption, too. Partial results (diagrams with several disconnected quark loops still need to be considered) from lattice simulations are promising, with small (about 5%) statistical uncertainty [213]. Various sources of systematic uncertainties are currently being investigated. For the fits, we take the result from Ref. 210, shifted by  $2 \times 10^{-11}$  to account for the more accurate charm quark treatment of Ref. 208, and with increased error to cover all recent evaluations, resulting in  $a_\mu^{\text{LBS}}(\alpha^3) = (+1.07 \pm 0.32) \times 10^{-9}$ .

Other hadronic effects at three-loop order [214] (updated in Ref. 20) contribute  $a_\mu^{\text{had}}(\alpha^3) = (-0.99 \pm 0.01) \times 10^{-9}$ . Correlations with the two-loop hadronic contribution and with  $\Delta\alpha(M_Z)$  (see Sec. 10.2) were considered in Ref. 199 which also contains analytic results for the perturbative QCD contribution. Very recently, hadronic four-loop effects have also been obtained, where the contributions without hadronic LBS subgraphs [215] amount to  $a_\mu^{\text{had}}(\alpha^4) = (0.123 \pm 0.001) \times 10^{-9}$  [20]. The contributions with a hadronic LBS subgraph have been estimated in Ref. 216, with the result  $a_\mu^{\text{LBS}}(\alpha^4) = (0.03 \pm 0.02) \times 10^{-9}$ .

Altogether, the SM prediction is

$$a_\mu^{\text{theory}} = (1165917.63 \pm 0.46) \times 10^{-9}, \quad (10.55)$$

<sup>†</sup> In what follows, we summarize the most important aspects of  $g_\mu - 2$ , and give some details on the evaluation in our fits. For more details see the dedicated contribution on “The Muon Anomalous Magnetic Moment” in this *Review*. There are some numerical differences, which are well understood and arise because internal consistency of the fits requires the calculation of all observables from analytical expressions and common inputs and fit parameters, so that an independent evaluation is necessary for this Section. Note, that in the spirit of a global analysis based on all available information we have chosen here to use an analysis [20] which considers the  $\tau$  decay data, as well.



where the error is from the hadronic uncertainties excluding parametric ones such as from  $\alpha_s$  and the heavy quark masses. Estimating a correlation of about 68% from the data input to the vacuum polarization integrals, we evaluate the correlation of the total (experimental plus theoretical) uncertainty in  $a_\mu$  with  $\Delta\alpha(M_Z)$  as 24%. The overall  $4.2\sigma$  discrepancy between the experimental and theoretical  $a_\mu$  values could be due to fluctuations (the E821 result is statistics dominated) or underestimates of the theoretical uncertainties. On the other hand, the deviation could also arise from physics beyond the SM, such as supersymmetric models with large  $\tan\beta$  and moderately light superparticle masses [217], or a dark  $Z$  boson [186].

## 10.6. Global fit results

In this section we present the results of global fits to the experimental data discussed in Sec. 10.3–Sec. 10.5. For earlier analyses see Refs. [10,117,218]

**Table 10.4:** Principal non- $Z$  pole observables, compared with the SM best fit predictions. The first  $M_W$  and  $\Gamma_W$  values are from the Tevatron [219,220] and the second ones from LEP 2 [173]. The value of  $m_t$  differs from the one in the Particle Listings since it includes recent preliminary results. The world averages for  $g_{V,A}^{\nu e}$  are dominated by the CHARM II [86] results,  $g_V^{\nu e} = -0.035 \pm 0.017$  and  $g_A^{\nu e} = -0.503 \pm 0.017$ . The errors are the total (experimental plus theoretical) uncertainties. The  $\tau_\tau$  value is the  $\tau$  lifetime world average computed by combining the direct measurements with values derived from the leptonic branching ratios [45]; in this case, the theory uncertainty is included in the SM prediction. In all other SM predictions, the uncertainty is from  $M_Z$ ,  $M_H$ ,  $m_t$ ,  $m_b$ ,  $m_c$ ,  $\hat{\alpha}(M_Z)$ , and  $\alpha_s$ , and their correlations have been accounted for. The column denoted Pull gives the standard deviations.

Quantity	Value	Standard Model	Pull
$m_t$ [GeV]	$173.34 \pm 0.81$	$173.76 \pm 0.76$	-0.5
$M_W$ [GeV]	$80.387 \pm 0.016$ $80.376 \pm 0.033$	$80.361 \pm 0.006$	1.6 0.4
$\Gamma_W$ [GeV]	$2.046 \pm 0.049$ $2.195 \pm 0.083$	$2.089 \pm 0.001$	-0.9 1.3
$M_H$ [GeV]	$125.09 \pm 0.24$	$125.11 \pm 0.24$	0.0
$\rho_{\gamma W}$	$-0.03 \pm 0.20$	$-0.02 \pm 0.02$	0.0
$\rho_{\tau Z}$	$-0.27 \pm 0.31$	$0.00 \pm 0.03$	-0.9
$g_V^{\nu e}$	$-0.040 \pm 0.015$	$-0.0397 \pm 0.0002$	0.0
$g_A^{\nu e}$	$-0.507 \pm 0.014$	$-0.5064$	0.0
$Q_W(e)$	$-0.0403 \pm 0.0053$	$-0.0473 \pm 0.0003$	1.3
$Q_W(p)$	$0.064 \pm 0.012$	$0.0708 \pm 0.0003$	-0.6
$Q_W(\text{Cs})$	$-72.62 \pm 0.43$	$-73.25 \pm 0.02$	1.5
$Q_W(\text{Tl})$	$-116.4 \pm 3.6$	$-116.91 \pm 0.02$	0.1
$s_Z^2(\text{eDIS})$	$0.2299 \pm 0.0043$	$0.23129 \pm 0.00005$	-0.3
$\tau_\tau$ [fs]	$290.88 \pm 0.35$	$289.85 \pm 2.12$	0.4
$\frac{1}{2}(g_\mu - 2 - \frac{\alpha}{\pi})$	$(4511.18 \pm 0.78) \times 10^{-9}$	$(4507.89 \pm 0.08) \times 10^{-9}$	4.2

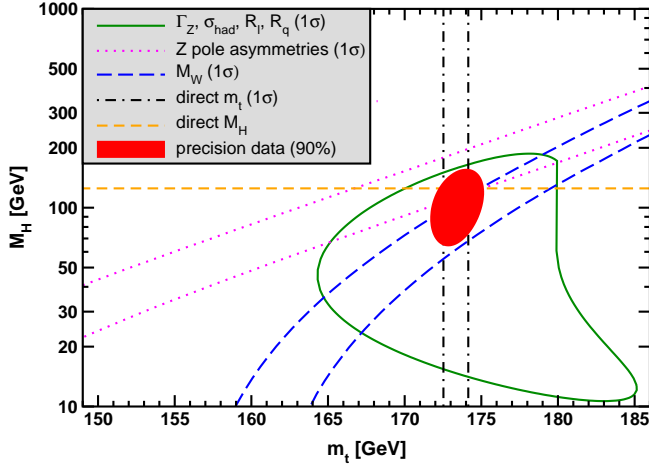
The values for  $m_t$  [48],  $M_W$  [173,219],  $\Gamma_W$  [173,220],  $M_H$  and the ratios of Higgs branching fractions [184] discussed in Sec. 10.4.5,  $\nu$ -lepton scattering [83–88], the weak charges of the electron [121], the proton [126], cesium [129,130] and thallium [131], the weak mixing angle extracted from eDIS [113], the muon anomalous magnetic moment [196], and the  $\tau$  lifetime are listed in Table 10.4. Likewise, the principal  $Z$  pole observables can be found in Table 10.5 where the LEP 1 averages of the ALEPH, DELPHI, L3 and OPAL results include common systematic errors and correlations [10]. The heavy flavor results of LEP 1 and SLD are based on common inputs and correlated, as well [10]. Note that the values of  $\Gamma(\ell^+\ell^-)$ ,  $\Gamma(\text{had})$ , and  $\Gamma(\text{inv})$  are not independent of  $\Gamma_Z$ , the  $R_\ell$ , and  $\sigma_{\text{had}}$  and that the SM errors in those latter are largely dominated by the uncertainty in

**Table 10.5:** Principal  $Z$  pole observables and their SM predictions (*cf.* Table 10.4). The first  $\bar{s}_\ell^2$  is the effective weak mixing angle extracted from the hadronic charge asymmetry, the second is the combined value from the Tevatron [164–166], and the third from the LHC [170–172]. The values of  $A_e$  are (i) from  $A_{LR}$  for hadronic final states [159]; (ii) from  $A_{LR}$  for leptonic final states and from polarized Bhabha scattering [161]; and (iii) from the angular distribution of the  $\tau$  polarization at LEP 1. The  $A_\tau$  values are from SLD and the total  $\tau$  polarization, respectively.

Quantity	Value	Standard Model	Pull
$M_Z$ [GeV]	$91.1876 \pm 0.0021$	$91.1880 \pm 0.0020$	-0.2
$\Gamma_Z$ [GeV]	$2.4952 \pm 0.0023$	$2.4943 \pm 0.0008$	0.4
$\Gamma(\text{had})$ [GeV]	$1.7444 \pm 0.0020$	$1.7420 \pm 0.0008$	—
$\Gamma(\text{inv})$ [MeV]	$499.0 \pm 1.5$	$501.66 \pm 0.05$	—
$\Gamma(\ell^+\ell^-)$ [MeV]	$83.984 \pm 0.086$	$83.995 \pm 0.010$	—
$\sigma_{\text{had}}[\text{nb}]$	$41.541 \pm 0.037$	$41.484 \pm 0.008$	1.5
$R_e$	$20.804 \pm 0.050$	$20.734 \pm 0.010$	1.4
$R_\mu$	$20.785 \pm 0.033$	$20.734 \pm 0.010$	1.6
$R_\tau$	$20.764 \pm 0.045$	$20.779 \pm 0.010$	-0.3
$R_b$	$0.21629 \pm 0.00066$	$0.21579 \pm 0.00003$	0.8
$R_c$	$0.1721 \pm 0.0030$	$0.17221 \pm 0.00003$	0.0
$A_{FB}^{(0,e)}$	$0.0145 \pm 0.0025$	$0.01622 \pm 0.00009$	-0.7
$A_{FB}^{(0,\mu)}$	$0.0169 \pm 0.0013$		0.5
$A_{FB}^{(0,\tau)}$	$0.0188 \pm 0.0017$		1.5
$A_{FB}^{(0,b)}$	$0.0992 \pm 0.0016$	$0.1031 \pm 0.0003$	-2.4
$A_{FB}^{(0,c)}$	$0.0707 \pm 0.0035$	$0.0736 \pm 0.0002$	-0.8
$A_{FB}^{(0,s)}$	$0.0976 \pm 0.0114$	$0.1032 \pm 0.0003$	-0.5
$\bar{s}_\ell^2$	$0.2324 \pm 0.0012$ $0.23185 \pm 0.00035$ $0.23105 \pm 0.00087$	$0.23152 \pm 0.00005$	0.7 0.9 -0.5
$A_e$	$0.15138 \pm 0.00216$ $0.1544 \pm 0.0060$ $0.1498 \pm 0.0049$	$0.1470 \pm 0.0004$	2.0 1.2 0.6
$A_\mu$	$0.142 \pm 0.015$		-0.3
$A_\tau$	$0.136 \pm 0.015$ $0.1439 \pm 0.0043$		-0.7 -0.7
$A_b$	$0.923 \pm 0.020$	$0.9347$	-0.6
$A_c$	$0.670 \pm 0.027$	$0.6678 \pm 0.0002$	0.1
$A_s$	$0.895 \pm 0.091$	$0.9356$	-0.4

$\alpha_s$ . Also shown in both tables are the SM predictions for the values of  $M_Z$ ,  $M_H$ ,  $\alpha_s(M_Z)$ ,  $\Delta\alpha_{\text{had}}^{(3)}$  and the heavy quark masses shown in Table 10.6. The predictions result from a global least-square ( $\chi^2$ ) fit to all data using the minimization package MINUIT [221] and the EW library GAPP [21]. In most cases, we treat all input errors (the uncertainties of the values) as Gaussian. The reason is not that we assume that theoretical and systematic errors are intrinsically bell-shaped (which they are not) but because in most cases the input errors are either dominated by the statistical components or they are combinations of many different (including statistical) error sources, which should yield approximately Gaussian combined errors by the large number theorem. An exception is the theory dominated error on the  $\tau$  lifetime, which we recalculate in each  $\chi^2$ -function call since it depends itself on  $\alpha_s$ . Sizes and shapes of the output errors (the uncertainties of the predictions and the SM fit parameters) are fully determined by the fit, and  $1\sigma$  errors are defined to correspond to  $\Delta\chi^2 = \chi^2 - \chi_{\text{min}}^2 = 1$ , and do not necessarily correspond to the 68.3% probability range or the 39.3% probability contour (for 2 parameters).

The agreement is generally very good. Despite the few discrepancies discussed in the following, the fit describes the data well, with a  $\chi^2/\text{d.o.f.} = 53.6/42$ . The probability of a larger  $\chi^2$  is 11%. Only



**Figure 10.4:** Fit result and one-standard-deviation (39.35% for the closed contours and 68% for the others) uncertainties in  $M_H$  as a function of  $m_t$  for various inputs, and the 90% CL region ( $\Delta\chi^2 = 4.605$ ) allowed by all data.  $\alpha_s(M_Z) = 0.1182$  is assumed except for the fits including the  $Z$  lineshape. The width of the horizontal dashed (yellow) band is not visible on the scale of the plot.

the final result for  $g_\mu - 2$  from BNL is currently showing a large ( $4.2\sigma$ ) conflict. In addition,  $A_{FB}^{(0,b)}$  from LEP 1 and  $A_{LR}^0$  (SLD) from hadronic final states deviate at the  $2\sigma$  level.  $g_L^2$  from NuTeV is nominally in conflict with the SM, as well, but the precise status is under investigation (see Sec. 10.3).

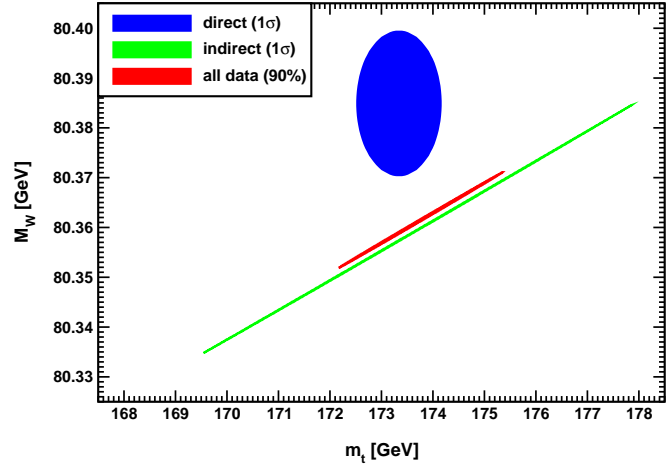
$A_b$  can be extracted from  $A_{FB}^{(0,b)}$  when  $A_e = 0.1501 \pm 0.0016$  is taken from a fit to leptonic asymmetries (using lepton universality). The result,  $A_b = 0.881 \pm 0.017$ , is  $3.1\sigma$  below the SM prediction<sup>§</sup> and also  $1.6\sigma$  below  $A_b = 0.923 \pm 0.020$  obtained from  $A_{LR}^{FB}(b)$  at SLD. Thus, it appears that at least some of the problem in  $A_b$  is due to a statistical fluctuation or other experimental effect in one of the asymmetries. Note, however, that the uncertainty in  $A_{FB}^{(0,b)}$  is strongly statistics dominated. The combined value,  $A_b = 0.899 \pm 0.013$  deviates by  $2.8\sigma$ .

The left-right asymmetry,  $A_{LR}^0 = 0.15138 \pm 0.00216$  [159], based on all hadronic data from 1992–1998 differs  $2.0\sigma$  from the SM expectation of  $0.1470 \pm 0.0004$ . The combined value of  $A_\ell = 0.1513 \pm 0.0021$  from SLD (using lepton-family universality and including correlations) is also  $2.0\sigma$  above the SM prediction; but there is experimental agreement between this SLD value and the LEP 1 value,  $A_\ell = 0.1481 \pm 0.0027$ , obtained from a fit to  $A_{FB}^{(0,\ell)}$ ,  $A_e(\mathcal{P}_\tau)$ , and  $A_\tau(\mathcal{P}_\tau)$ , again assuming universality.

The observables in Table 10.4 and Table 10.5, as well as some other less precise observables, are used in the global fits described below. In all fits, the errors include full statistical, systematic, and theoretical uncertainties. The correlations on the LEP 1 lineshape and  $\tau$  polarization, the LEP/SLD heavy flavor observables, the SLD lepton asymmetries, and the  $\nu$ - $e$  scattering observables, are included. The theoretical correlations between  $\Delta\alpha_{\text{had}}^{(5)}$  and  $g_\mu - 2$ , and between the charm and bottom quark masses, are also accounted for.

The electroweak data allow a simultaneous determination of  $M_Z$ ,  $M_H$ ,  $m_t$ , and the strong coupling  $\alpha_s(M_Z)$ . ( $\hat{m}_c$ ,  $\hat{m}_b$ , and  $\Delta\alpha_{\text{had}}^{(3)}$  are also allowed to float in the fits, subject to the theoretical constraints [19,45] described in Sec. 10.2. These are correlated with  $\alpha_s$ .)  $\alpha_s$  is determined mainly from  $R_\ell$ ,  $\Gamma_Z$ ,  $\sigma_{\text{had}}$ , and  $\tau_\tau$ . The global fit to all data, including the hadron collider average  $m_t = 173.34 \pm 0.81$  GeV, yields the result in Table 10.6 (the  $\overline{\text{MS}}$  top quark mass given there corresponds to  $m_t = 173.76 \pm 0.76$  GeV). The

<sup>§</sup> Alternatively, one can use  $A_\ell = 0.1481 \pm 0.0027$ , which is from LEP 1 alone and in excellent agreement with the SM, and obtain  $A_b = 0.893 \pm 0.022$  which is  $1.9\sigma$  low. This illustrates that some of the discrepancy is related to the one in  $A_{LR}$ .



**Figure 10.5:** One-standard-deviation (39.35%) region in  $M_W$  as a function of  $m_t$  for the direct and indirect data, and the 90% CL region ( $\Delta\chi^2 = 4.605$ ) allowed by all data.

weak mixing angle, see Table 10.2, is determined to

$$\hat{s}_Z^2 = 0.23129 \pm 0.00005, \quad \hat{s}_W^2 = 0.22336 \pm 0.00010,$$

while the corresponding effective angle is  $\hat{s}_\ell^2 = 0.23152 \pm 0.00005$ .

As a cross-check, one can also perform a fit without the direct mass constraint,  $M_H = 125.09 \pm 0.24$  GeV, in Eq. (10.49). In this case we obtain a 2% indirect mass determination,

$$M_H = 126.1 \pm 1.9 \text{ GeV}, \quad (10.56)$$

arising predominantly from the quantities in Eq. (10.51), since the branching ratio for  $H \rightarrow ZZ^*$  varies very rapidly as a function of  $M_H$  for Higgs masses near 125 GeV. Removing also the branching ratio constraints gives the loop-level determination from the precision data alone,

$$M_H = 96_{-19}^{+22} \text{ GeV}, \quad (10.57)$$

which is  $1.2\sigma$  below the kinematical constraint, but the latter is inside the 90% central confidence range,

$$66 \text{ GeV} < M_H < 134 \text{ GeV}. \quad (10.58)$$

This is mostly a reflection of the Tevatron determination of  $M_W$ , which is  $1.6\sigma$  higher than the SM best fit value in Table 10.4. This is illustrated in Fig. 10.4 where one sees that the precision data together with  $M_H$  from the LHC prefer that  $m_t$  is closer to the upper end of its  $1\sigma$  allowed range. Conversely, one can remove the direct  $M_W$  and  $\Gamma_W$  constraints from the fits and use Eq. (10.49) to obtain  $M_W = 80.357 \pm 0.006$  GeV. This is  $1.7\sigma$  below the Tevatron/LEP 2 average,  $M_W = 80.385 \pm 0.015$  GeV.

Finally, one can carry out a fit without including the constraint,  $m_t = 173.34 \pm 0.81$  GeV, from the hadron colliders. (The indirect prediction is for the  $\overline{\text{MS}}$  mass,  $\hat{m}_t(\hat{m}_t) = 166.8 \pm 2.0$  GeV, which is in the end converted to the pole mass.) One obtains  $m_t = 176.7 \pm 2.1$  GeV, which is  $1.5\sigma$  higher than the direct Tevatron/LHC average. The situation is summarized in Fig. 10.5 showing the  $1\sigma$  contours in the  $M_W$ - $m_t$  plane from the direct and indirect determinations, as well as the combined 90% CL region.

As described in the paragraph following Eq. (10.54) in Sec. 10.5, there is some stress in the experimental  $e^+e^-$  and  $\tau$  spectral functions. These are below or above the  $2\sigma$  level (depending on what is actually compared) but not larger than the deviations of some other quantities entering our analyses. The number and size of these deviations are not inconsistent with what one would expect to happen as a result of random fluctuations. It is nevertheless instructive to study the effect of doubling the uncertainty in  $\Delta\alpha_{\text{had}}^{(3)}$  (2 GeV)  $= (58.04 \pm 1.10) \times 10^{-4}$  (see Sec. 10.2) on the loop-level determination. The result,  $M_H = 90_{-19}^{+23}$  GeV, deviates even

slightly *more* ( $1.4\sigma$ ) than Eq. (10.57), and demonstrates that the uncertainty in  $\Delta\alpha_{\text{had}}$  is currently of only secondary importance. Note also that a shift of  $\pm 10^{-4}$  in  $\Delta\alpha_{\text{had}}^{(3)}(2\text{ GeV})$  corresponds to a shift of  $\mp 4.5\text{ GeV}$  in  $M_H$ . The hadronic contribution to  $\alpha(M_Z)$  is correlated with  $g_\mu - 2$  (see Sec. 10.5). The measurement of the latter is higher than the SM prediction, and its inclusion in the fit favors a larger  $\alpha(M_Z)$  and a lower  $M_H$  from the precision data (currently by  $6.4\text{ GeV}$ ).

of the SM. One should keep in mind, however, that this value,  $\alpha_s(M_Z) = 0.1203 \pm 0.0028$ , is very sensitive to certain types of new physics such as non-universal vertex corrections. In contrast, the value derived from  $\tau$  decays,  $\alpha_s(M_Z) = 0.1174^{+0.0019}_{-0.0017}$ , is theory dominated but less sensitive to new physics. The two values are in reasonable agreement with each other. They are also in good agreement with the averages from jet-event shapes in  $e^+e^-$  annihilation ( $0.1169 \pm 0.0034$ ) and lattice simulations ( $0.1187 \pm 0.0012$ ), whereas the DIS average

**Table 10.6:** Principal SM fit result including mutual correlations (all masses in GeV).

$M_Z$	$91.1880 \pm 0.0020$	1.00	-0.08	-0.01	-0.02	0.01	0.04	0.00
$\hat{m}_t(\hat{m}_t)$	$164.08 \pm 0.73$	-0.08	1.00	-0.01	-0.08	-0.17	0.07	-0.01
$\hat{m}_b(\hat{m}_b)$	$4.199 \pm 0.023$	-0.01	-0.01	1.00	0.23	-0.04	0.02	0.00
$\hat{m}_c(\hat{m}_c)$	$1.265^{+0.030}_{-0.038}$	-0.02	-0.08	0.23	1.00	0.10	0.07	0.00
$\alpha_s(M_Z)$	$0.1182 \pm 0.0016$	0.01	-0.17	-0.04	0.10	1.00	-0.04	-0.01
$\Delta\alpha_{\text{had}}^{(3)}(2\text{ GeV})$	$0.00590 \pm 0.00011$	0.04	0.07	0.02	0.07	-0.04	1.00	0.00
$M_H$	$125.11 \pm 0.24$	0.00	-0.01	0.00	0.00	-0.01	0.00	1.00

**Table 10.7:** Values of  $\hat{s}_Z^2$ ,  $s_W^2$ ,  $\alpha_s$ ,  $m_t$  and  $M_H$  [both in GeV] for various data sets. The  $M_H$  constraint refers collectively to the kinematical and decay information from Sec. 10.4.5. In the fit to the LHC (Tevatron) data the  $\alpha_s$  constraint is from the  $t\bar{t}$  production [222] (inclusive jet [223]) cross-section. The  $m_t$  input for the LHC is taken from Ref. 224.

Data	$\hat{s}_Z^2$	$s_W^2$	$\alpha_s(M_Z)$	$m_t$	$M_H$
All data	0.23129(5)	0.22336(10)	0.1182(16)	$173.8 \pm 0.8$	$125.1 \pm 0.2$
All data except $M_H$	0.23119(10)	0.22312(22)	0.1184(16)	$173.4 \pm 0.8$	$96^{+22}_{-19}$
All data except $M_Z$	0.23122(7)	0.22332(11)	0.1181(16)	$173.4 \pm 0.8$	$125.1 \pm 0.2$
All data except $M_W$	0.23132(6)	0.22343(11)	0.1185(16)	$173.4 \pm 0.8$	$125.1 \pm 0.2$
All data except $m_t$	0.23122(7)	0.22303(24)	0.1185(16)	$176.7 \pm 2.1$	$125.1 \pm 0.2$
$M_H, M_Z, \Gamma_Z, m_t$	0.23129(9)	0.22342(16)	0.1202(45)	$173.3 \pm 0.8$	$125.1 \pm 0.2$
LHC	0.23081(88)	0.22298(88)	0.1151(27)	$172.4 \pm 0.8$	$125.1 \pm 0.2$
Tevatron + $M_Z$	0.23113(13)	0.22307(30)	0.1161(45)	$173.4 \pm 0.8$	$100^{+32}_{-26}$
LEP	0.23147(17)	0.22346(47)	0.1220(31)	$182 \pm 12$	$283^{+395}_{-158}$
SLD + $M_Z, \Gamma_Z, m_t$	0.23074(28)	0.22229(55)	0.1174(45)	$173.3 \pm 0.8$	$43^{+33}_{-23}$
$A_{FB}^{(b,c)}, M_Z, \Gamma_Z, m_t$	0.23200(29)	0.22508(70)	0.1277(51)	$173.3 \pm 0.8$	$380^{+219}_{-138}$
$M_{W,Z}, \Gamma_{W,Z}, m_t$	0.23106(14)	0.22292(29)	0.1185(43)	$173.3 \pm 0.8$	$83^{+26}_{-22}$
low energy + $M_{H,Z}$	0.2328(14)	0.2291(55)	0.1175(18)	$121 \pm 50$	$125.1 \pm 0.2$

The weak mixing angle can be determined from  $Z$  pole observables,  $M_W$ , and from a variety of neutral-current processes spanning a very wide  $Q^2$  range. The results (for the older low energy neutral-current data see Refs. 117 and 218, as well as earlier editions of this *Review*) shown in Table 10.7 are in reasonable agreement with each other, indicating the quantitative success of the SM. The largest discrepancy is the value  $\hat{s}_Z^2 = 0.23200 \pm 0.00029$  from the forward-backward asymmetries into bottom and charm quarks, which is  $2.4\sigma$  above the value  $0.23129 \pm 0.00005$  from the global fit to all data (see Table 10.5). Similarly,  $\hat{s}_Z^2 = 0.23074 \pm 0.00028$  from the SLD asymmetries (in both cases when combined with  $M_Z$ ) is  $2.0\sigma$  low.

The extracted  $Z$  pole value of  $\alpha_s(M_Z)$  is based on a formula with negligible theoretical uncertainty if one assumes the exact validity

( $0.1156 \pm 0.0023$ ) is somewhat lower than the  $Z$  pole value. For more details, other determinations, and references, see Section 9 on “Quantum Chromodynamics” in this *Review*.

Using  $\alpha(M_Z)$  and  $\hat{s}_Z^2$  as inputs, one can predict  $\alpha_s(M_Z)$  assuming grand unification. One finds [225]  $\alpha_s(M_Z) = 0.130 \pm 0.001 \pm 0.01$  for the simplest theories based on the minimal supersymmetric extension of the SM, where the first (second) uncertainty is from the inputs (thresholds). This is slightly larger, but consistent with  $\alpha_s(M_Z) = 0.1182 \pm 0.0016$  from our fit, as well as with most other determinations. Non-supersymmetric unified theories predict the low value  $\alpha_s(M_Z) = 0.073 \pm 0.001 \pm 0.001$ . See also the note on “Supersymmetry” in the Searches Particle Listings.

Most of the parameters relevant to  $\nu$ -hadron,  $\nu$ - $e$ ,  $e$ -hadron,

and  $e^-e^\pm$  processes are determined uniquely and precisely from the data in “model-independent” fits (*i.e.*, fits which allow for an arbitrary EW gauge theory). The values for the parameters defined in Eqs. (10.16)–(10.17) are given in Table 10.8 along with the predictions of the SM. The agreement is very good. (The  $\nu$ -hadron results including the original NuTeV data can be found in the 2006 edition of this *Review*, and fits with modified NuTeV constraints in the 2008 and 2010 editions.) The off  $Z$  pole  $e^+e^-$  results are difficult to present in a model-independent way because  $Z$  propagator effects are non-negligible at TRISTAN, PETRA, PEP, and LEP 2 energies. However, assuming  $e$ - $\mu$ - $\tau$  universality, the low energy lepton asymmetries imply [156]  $4(g_{AV}^e)^2 = 0.99 \pm 0.05$ , in good agreement with the SM prediction  $\simeq 1$ .

**Table 10.8:** Values of the model-independent neutral-current parameters, compared with the SM predictions. There is a second  $g_{LV,LA}^{\nu e}$  solution, given approximately by  $g_{LV}^{\nu e} \leftrightarrow g_{LA}^{\nu e}$ , which is eliminated by  $e^+e^-$  data under the assumption that the neutral current is dominated by the exchange of a single  $Z$  boson. The  $g_{LL}^{\nu q}$ , as well as the  $g_{LR}^{\nu q}$ , are strongly correlated and non-Gaussian, so that for implementations we recommend the parametrization using  $g_i^2$  and  $\tan \theta_i = g_{Li}^{\nu u}/g_{Li}^{\nu d}$  where  $i = L, R$ . In the SM predictions, the parametric uncertainties from  $M_Z$ ,  $M_H$ ,  $m_t$ ,  $m_b$ ,  $m_c$ ,  $\hat{\alpha}(M_Z)$ , and  $\alpha_s$  are negligible.

Quantity	Experimental Value	Standard Model	Correlation
$g_{LL}^{\nu u}$	$0.328 \pm 0.016$	0.3457	non-Gaussian
$g_{LL}^{\nu d}$	$-0.440 \pm 0.011$	-0.4288	
$g_{LR}^{\nu u}$	$-0.179 \pm 0.013$	-0.1553	
$g_{LR}^{\nu d}$	$-0.027^{+0.077}_{-0.048}$	0.0777	
$g_L^2$	$0.3005 \pm 0.0028$	0.3034	small
$g_R^2$	$0.0329 \pm 0.0030$	0.0302	
$\theta_L$	$2.50 \pm 0.035$	2.4631	
$\theta_R$	$4.56^{+0.42}_{-0.27}$	5.1765	
$g_{LV}^{\nu e}$	$-0.040 \pm 0.015$	-0.0396	-0.05
$g_{LA}^{\nu e}$	$-0.507 \pm 0.014$	-0.5064	
$g_{AV}^{eu} + 2g_{AV}^{ed}$	$0.489 \pm 0.005$	0.4951	-0.94 0.42
$2g_{AV}^{eu} - g_{AV}^{ed}$	$-0.708 \pm 0.016$	-0.7192	-0.45
$2g_{VA}^{eu} - g_{VA}^{ed}$	$-0.144 \pm 0.068$	-0.0949	
$g_{VA}^{ee}$	$0.0190 \pm 0.0027$	0.0225	

## 10.7. Constraints on new physics

The masses and decay properties of the electroweak bosons and low energy data can be used to search for and set limits on deviations from the SM. We will mainly discuss the effects of exotic particles (with heavy masses  $M_{\text{new}} \gg M_Z$  in an expansion in  $M_Z/M_{\text{new}}$ ) on the gauge boson self-energies. (Brief remarks are made on new physics which is not of this type.) Most of the effects on precision measurements can be described by three gauge self-energy parameters  $S$ ,  $T$ , and  $U$ . We will define these, as well as the related parameters  $\rho_0$ ,  $\epsilon_i$ , and  $\hat{\epsilon}_i$ , to arise from new physics only. In other words, they are equal to zero ( $\rho_0 = 1$ ) exactly in the SM, and do not include any (loop induced) contributions that depend on  $m_t$  or  $M_H$ , which are treated separately. Our treatment differs from most of the original papers.

The dominant effect of many extensions of the SM can be described by the  $\rho_0$  parameter,

$$\rho_0 \equiv \frac{M_W^2}{M_Z^2 \hat{c}_Z^2 \hat{\rho}}, \quad (10.59)$$

which describes new sources of SU(2) breaking that cannot be accounted for by the SM Higgs doublet or  $m_t$  effects.  $\hat{\rho}$  is calculated as in Eq. (10.12) assuming the validity of the SM. In the presence

of  $\rho_0 \neq 1$ , Eq. (10.59) generalizes the second Eq. (10.12) while the first remains unchanged. Provided that the new physics which yields  $\rho_0 \neq 1$  is a small perturbation which does not significantly affect other radiative corrections,  $\rho_0$  can be regarded as a phenomenological parameter which multiplies  $G_F$  in Eqs. (10.16)–(10.17), (10.32), and  $\Gamma_Z$  in Eq. (10.46c). There are enough data to determine  $\rho_0$ ,  $M_H$ ,  $m_t$ , and  $\alpha_s$ , simultaneously. From the global fit,

$$\rho_0 = 1.00037 \pm 0.00023, \quad (10.60)$$

$$\alpha_s(M_Z) = 0.1183 \pm 0.0016, \quad (10.61)$$

The result in Eq. (10.60) is 1.6  $\sigma$  above the SM expectation,  $\rho_0 = 1$ . It can be used to constrain higher-dimensional Higgs representations to have vacuum expectation values of less than a few percent of those of the doublets. Indeed, the relation between  $M_W$  and  $M_Z$  is modified if there are Higgs multiplets with weak isospin  $> 1/2$  with significant vacuum expectation values. For a general (charge-conserving) Higgs structure,

$$\rho_0 = \frac{\sum_i [t(i)(t(i)+1) - t_3(i)^2] |v_i|^2}{2 \sum_i t_3(i)^2 |v_i|^2}, \quad (10.62)$$

where  $v_i$  is the expectation value of the neutral component of a Higgs multiplet with weak isospin  $t(i)$  and third component  $t_3(i)$ . In order to calculate to higher orders in such theories one must define a set of four fundamental renormalized parameters which one may conveniently choose to be  $\alpha$ ,  $G_F$ ,  $M_Z$ , and  $M_W$ , since  $M_W$  and  $M_Z$  are directly measurable. Then  $\hat{s}_Z^2$  and  $\rho_0$  can be considered dependent parameters.

Eq. (10.60) can also be used to constrain other types of new physics. For example, non-degenerate multiplets of heavy fermions or scalars break the vector part of weak SU(2) and lead to a decrease in the value of  $M_Z/M_W$ . Each non-degenerate SU(2) doublet  $(\frac{f_1}{f_2})$  yields a positive contribution to  $\rho_0$  [226] of

$$\frac{C G_F}{8\sqrt{2}\pi^2} \Delta m^2, \quad (10.63)$$

where

$$\Delta m^2 \equiv m_1^2 + m_2^2 - \frac{4m_1^2 m_2^2}{m_1^2 - m_2^2} \ln \frac{m_1}{m_2} \geq (m_1 - m_2)^2, \quad (10.64)$$

and  $C = 1$  (3) for color singlets (triplets). Eq. (10.60) taken together with Eq. (10.63) implies the following constraint on the mass splitting at the 95% CL,

$$\sum_i \frac{C_i}{3} \Delta m_i^2 \leq (49 \text{ GeV})^2. \quad (10.65)$$

where the sum runs over all new-physics doublets, for example fourth-family quarks or leptons,  $(\frac{b'}{b'})$  or  $(\frac{\nu'}{\nu'})$ , vector-like fermion doublets (which contribute to the sum in Eq. (10.65) with an extra factor of 2), and scalar doublets such as  $(\frac{\tilde{t}}{\tilde{b}})$  in Supersymmetry (in the absence of  $L$ - $R$  mixing).

Non-degenerate multiplets usually imply  $\rho_0 > 1$ . Similarly, heavy  $Z'$  bosons decrease the prediction for  $M_Z$  due to mixing and generally lead to  $\rho_0 > 1$  [227]. On the other hand, additional Higgs doublets which participate in spontaneous symmetry breaking [228] or heavy lepton doublets involving Majorana neutrinos [229], both of which have more complicated expressions, as well as the vacuum expectation values of Higgs triplets or higher-dimensional representations can contribute to  $\rho_0$  with either sign. Allowing for the presence of heavy degenerate chiral multiplets (the  $S$  parameter, to be discussed below) affects the determination of  $\rho_0$  from the data, at present leading to a larger value.

A number of authors [230–235] have considered the general effects on neutral-current and  $Z$  and  $W$  boson observables of various types of heavy (*i.e.*,  $M_{\text{new}} \gg M_Z$ ) physics which contribute to the  $W$  and  $Z$  self-energies but which do not have any direct coupling to the ordinary fermions. In addition to non-degenerate multiplets, which break the vector part of weak SU(2), these include heavy degenerate multiplets of chiral fermions which break the axial generators.

Such effects can be described by just three parameters,  $S$ ,  $T$ , and  $U$ , at the (EW) one-loop level. (Three additional parameters are needed if the new physics scale is comparable to  $M_Z$  [236]. Further generalizations, including effects relevant to LEP 2, are described in Ref. 237.)  $T$  is proportional to the difference between the  $W$  and  $Z$  self-energies at  $Q^2 = 0$  (i.e., vector SU(2)-breaking), while  $S$  ( $S+U$ ) is associated with the difference between the  $Z$  ( $W$ ) self-energy at  $Q^2 = M_{Z,W}^2$  and  $Q^2 = 0$  (axial SU(2)-breaking). Denoting the contributions of new physics to the various self-energies by  $\Pi_{ij}^{\text{new}}$ , we have

$$\hat{\alpha}(M_Z)T \equiv \frac{\Pi_{WW}^{\text{new}}(0)}{M_W^2} - \frac{\Pi_{ZZ}^{\text{new}}(0)}{M_Z^2}, \quad (10.66a)$$

$$\frac{\hat{\alpha}(M_Z)}{4\hat{s}_Z^2\hat{c}_Z^2}S \equiv \frac{\Pi_{ZZ}^{\text{new}}(M_Z^2) - \Pi_{ZZ}^{\text{new}}(0)}{M_Z^2} - \frac{\hat{c}_Z^2 - \hat{s}_Z^2}{\hat{c}_Z\hat{s}_Z} \frac{\Pi_{Z\gamma}^{\text{new}}(M_Z^2)}{M_Z^2} - \frac{\Pi_{\gamma\gamma}^{\text{new}}(M_Z^2)}{M_Z^2}, \quad (10.66b)$$

$$\frac{\hat{\alpha}(M_Z)}{4\hat{s}_Z^2}(S+U) \equiv \frac{\Pi_{WW}^{\text{new}}(M_W^2) - \Pi_{WW}^{\text{new}}(0)}{M_W^2} - \frac{\hat{c}_Z}{\hat{s}_Z} \frac{\Pi_{Z\gamma}^{\text{new}}(M_Z^2)}{M_Z^2} - \frac{\Pi_{\gamma\gamma}^{\text{new}}(M_Z^2)}{M_Z^2}. \quad (10.66c)$$

$S$ ,  $T$ , and  $U$  are defined with a factor proportional to  $\hat{\alpha}$  removed, so that they are expected to be of order unity in the presence of new physics. In the  $\overline{\text{MS}}$  scheme as defined in Ref. 57, the last two terms in Eqs. (10.66b) and (10.66c) can be omitted (as was done in some earlier editions of this *Review*). These three parameters are related to other parameters ( $S_i$ ,  $h_i$ ,  $\hat{e}_i$ ) defined in Refs. [57,231,232] by

$$\begin{aligned} T &= h_V = \hat{e}_1/\hat{\alpha}(M_Z), \\ S &= h_{AZ} = S_Z = 4\hat{s}_Z^2\hat{e}_3/\hat{\alpha}(M_Z), \\ U &= h_{AW} - h_{AZ} = S_W - S_Z \\ &= -4\hat{s}_Z^2\hat{e}_2/\hat{\alpha}(M_Z). \end{aligned} \quad (10.67)$$

A heavy non-degenerate multiplet of fermions or scalars contributes positively to  $T$  as

$$\rho_0 - 1 = \frac{1}{1 - \hat{\alpha}(M_Z)T} - 1 \simeq \hat{\alpha}(M_Z)T, \quad (10.68)$$

where  $\rho_0 - 1$  is given in Eq. (10.63). The effects of non-standard Higgs representations cannot be separated from heavy non-degenerate multiplets unless the new physics has other consequences, such as vertex corrections. Most of the original papers defined  $T$  to include the effects of loops only. However, we will redefine  $T$  to include all new sources of SU(2) breaking, including non-standard Higgs, so that  $T$  and  $\rho_0$  are equivalent by Eq. (10.68).

A multiplet of heavy degenerate chiral fermions yields

$$S = \frac{C}{3\pi} \sum_i \left( t_{3L}(i) - t_{3R}(i) \right)^2, \quad (10.69)$$

where  $t_{3L,R}(i)$  is the third component of weak isospin of the left-(right-)handed component of fermion  $i$  and  $C$  is the number of colors. For example, a heavy degenerate ordinary or mirror family would contribute  $2/3\pi$  to  $S$ . In models with warped extra dimensions, sizeable correction to the  $S$  parameter are generated by mixing effects between the SM gauge bosons and their Kaluza-Klein (KK) excitations. One finds  $S \approx 30v^2/M_{KK}^2$ , where  $M_{KK}$  is the mass of the KK gauge bosons [238]. Large positive values  $S > 0$  can also be generated in Technicolor models with QCD-like dynamics, where one expects [230]  $S \sim 0.45$  for an iso-doublet of techni-fermions, assuming  $N_{TC} = 4$  techni-colors, while  $S \sim 1.62$  for a full techni-generation with  $N_{TC} = 4$ . However, the QCD-like models are excluded on other grounds (flavor changing neutral currents, too-light quarks and pseudo-Goldstone bosons [239], and absence of a Higgs-like scalar).

On the other hand, negative values  $S < 0$  are possible, for example, for models of walking Technicolor [240] or loops involving scalars

or Majorana particles [241]. The simplest origin of  $S < 0$  would probably be an additional heavy  $Z'$  boson [227]. Supersymmetric extensions of the SM generally give very small effects. See Refs. 242 and 243 and the note on ‘‘Supersymmetry’’ in the Searches Particle Listings for a complete set of references.

Most simple types of new physics yield  $U = 0$ , although there are counter-examples, such as the effects of anomalous triple gauge vertices [232].

The SM expressions for observables are replaced by

$$\begin{aligned} M_Z^2 &= M_{Z0}^2 \frac{1 - \hat{\alpha}(M_Z)T}{1 - G_F M_{Z0}^2 S/2\sqrt{2}\pi}, \\ M_W^2 &= M_{W0}^2 \frac{1}{1 - G_F M_{W0}^2 (S+U)/2\sqrt{2}\pi}, \end{aligned} \quad (10.70)$$

where  $M_{Z0}$  and  $M_{W0}$  are the SM expressions (as functions of  $m_t$  and  $M_H$ ) in the  $\overline{\text{MS}}$  scheme. Furthermore,

$$\Gamma_Z = \frac{M_Z^2 \beta_Z}{1 - \hat{\alpha}(M_Z)T}, \Gamma_W = M_W^3 \beta_W, A_i = \frac{A_{i0}}{1 - \hat{\alpha}(M_Z)T}, \quad (10.71)$$

where  $\beta_Z$  and  $\beta_W$  are the SM expressions for the reduced widths  $\Gamma_{Z0}/M_{Z0}^3$  and  $\Gamma_{W0}/M_{W0}^3$ ,  $M_Z$  and  $M_W$  are the physical masses, and  $A_i$  ( $A_{i0}$ ) is a neutral-current amplitude (in the SM).

The data allow a simultaneous determination of  $\hat{s}_Z^2$  (from the  $Z$  pole asymmetries),  $S$  (from  $M_Z$ ),  $U$  (from  $M_W$ ),  $T$  (mainly from  $\Gamma_Z$ ),  $\alpha_s$  (from  $R_\ell$ ,  $\sigma_{\text{had}}$ , and  $\tau_\tau$ ),  $M_H$  and  $m_t$  (from the hadron colliders), with little correlation among the SM parameters:

$$\begin{aligned} S &= 0.05 \pm 0.10, \\ T &= 0.08 \pm 0.12, \\ U &= 0.02 \pm 0.10, \end{aligned} \quad (10.72)$$

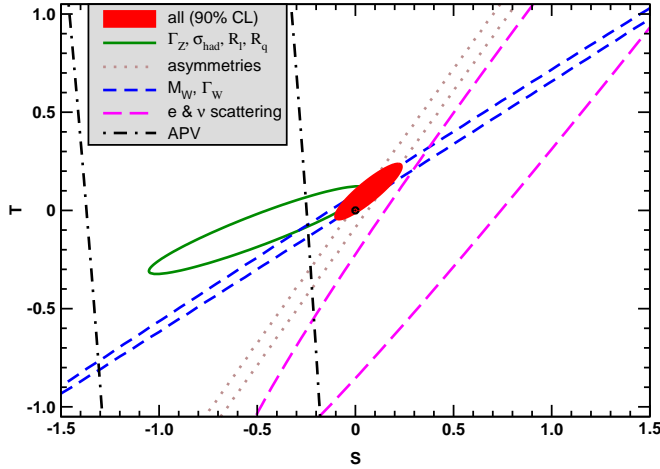
$\hat{s}_Z^2 = 0.23131 \pm 0.00015$ , and  $\alpha_s(M_Z) = 0.1182 \pm 0.0017$ , where the uncertainties are from the inputs. The parameters in Eqs. (10.72), which by definition are due to new physics only, are in excellent agreement with the SM values of zero. Fixing  $U = 0$  (as is also done in Fig. 10.6) moves  $S$  and  $T$  slightly upwards,

$$\begin{aligned} S &= 0.07 \pm 0.08, \\ T &= 0.10 \pm 0.07, \end{aligned} \quad (10.73)$$

with  $T$  showing a  $1.5\sigma$  deviation from zero. Using Eq. (10.68), the value of  $\rho_0$  corresponding to  $T$  in Eq. (10.72) is  $1.0006 \pm 0.0009$ , while the one corresponding to Eq. (10.73) is  $1.0008 \pm 0.0005$ .

There is a strong correlation (91%) between the  $S$  and  $T$  parameters. The  $U$  parameter is  $-61\%$  ( $-82\%$ ) anti-correlated with  $S$  ( $T$ ). The allowed regions in  $S$ - $T$  are shown in Fig. 10.6. From Eqs. (10.72) one obtains  $S < 0.22$  and  $T < 0.27$  at 95% CL, where the former puts the constraint  $M_{KK} \gtrsim 2.9$  TeV on the masses of KK gauge bosons in warped extra dimensions.

The  $S$  parameter can also be used to constrain the number of fermion families, *under the assumption* that there are no new contributions to  $T$  or  $U$  and therefore that any new families are degenerate; then an extra generation of SM fermions is excluded at the  $7\sigma$  level corresponding to  $N_F = 2.83 \pm 0.17$ . This can be compared to the fit to the number of light neutrinos given in Eq. (10.37),  $N_\nu = 2.992 \pm 0.007$ . However, the  $S$  parameter fits are valid even for a very heavy fourth family neutrino. Allowing  $T$  to vary as well, the constraint on a fourth family is weaker [244]. However, a heavy fourth family would increase the Higgs production cross-section through gluon fusion by a factor  $\sim 9$ , which is in considerable tension with the observed Higgs signal at LHC. Combining the limits from electroweak precision data with the measured Higgs production rate and limits from direct searches for heavy quarks [245], a fourth family of chiral fermions is now excluded by more than five standard deviations [246]. Similar remarks apply to a heavy mirror family [247] involving right-handed SU(2) doublets and left-handed singlets. In contrast, new doublets that receive most of their mass from a different



**Figure 10.6:**  $1\sigma$  constraints (39.35% for the closed contours and 68% for the others) on  $S$  and  $T$  (for  $U = 0$ ) from various inputs combined with  $M_Z$ .  $S$  and  $T$  represent the contributions of new physics only. Data sets not involving  $M_W$  or  $\Gamma_W$  are insensitive to  $U$ . With the exception of the fit to all data, we fix  $\alpha_s = 0.1182$ . The black dot indicates the Standard Model values  $S = T = 0$ .

source than the Higgs vacuum expectation value, such as vector-like fermion doublets or scalar doublets in Supersymmetry, give small or no contribution to  $S$ ,  $T$ ,  $U$  and the Higgs production cross-section and thus are still allowed. Partial or complete vector-like fermion families are predicted in many grand unified theories [248].

As discussed in Sec. 10.6, there is a 4.0% deviation in the asymmetry parameter  $A_b$ . Assuming that this is due to new physics affecting preferentially the third generation, we can perform a fit allowing additional  $Z \rightarrow b\bar{b}$  vertex corrections  $\rho_b$  and  $\kappa_b$  as in Eq. (10.33) (here defined to be due to new physics only with the SM contributions removed), as well as  $S$ ,  $T$ ,  $U$ , and the SM parameters, with the result,

$$\rho_b = 0.056 \pm 0.020, \quad \kappa_b = 0.182 \pm 0.068, \quad (10.74)$$

with an almost perfect correlation of 99% (because  $R_b$  is much better determined than  $A_b$ ). The central values of the oblique parameters are close to their SM values of zero, and there is little change in the SM parameters, except that the value of  $\alpha_s(M_Z)$  is lower by 0.0006 compared to the SM fit. Given that almost a  $\sim 20\%$  correction to  $\kappa_b$  would be necessary, it would be difficult to account for the deviation in  $A_b$  by new physics that enters only at the level of radiative corrections. Thus, if it is due to new physics, it is most likely of tree-level type affecting preferentially the third generation. Examples include the decay of a scalar neutrino resonance [249], mixing of the  $b$  quark with heavy exotics [250], and a heavy  $Z'$  with family non-universal couplings [251,252]. It is difficult, however, to simultaneously account for  $R_b$  without tuning, which has been measured on the  $Z$  peak and off-peak [253] at LEP 1. An average of  $R_b$  measurements at LEP 2 at energies between 133 and 207 GeV is  $2.1\sigma$  below the SM prediction, while  $A_{FB}^{(b)}$  (LEP 2) is  $1.6\sigma$  low [174].

There is no simple parametrization to describe the effects of every type of new physics on every possible observable. The  $S$ ,  $T$ , and  $U$  formalism describes many types of heavy physics which affect only the gauge self-energies, and it can be applied to all precision observables. However, new physics which couples directly to ordinary fermions, such as heavy  $Z'$  bosons [227], mixing with exotic fermions [254], or leptoquark exchange [173,255] cannot be fully parametrized in the  $S$ ,  $T$ , and  $U$  framework. It is convenient to treat these types of new physics by parameterizations that are specialized to that particular class of theories (e.g., extra  $Z'$  bosons), or to consider specific models (which might contain, e.g.,  $Z'$  bosons and exotic fermions with correlated parameters). Fits to Supersymmetric models are described in Ref. 243. Models involving strong dynamics (such as composite Higgs scenarios) for EW breaking are considered in Ref. 256. The

effects of compactified extra spatial dimensions at the TeV scale are reviewed in Ref. 257, and constraints on Little Higgs models in Ref. 258. The implications of non-standard Higgs sectors, e.g., involving Higgs singlets or triplets, are discussed in Ref. 259, while additional Higgs doublets are considered in Refs. 228 and 260. Limits on new four-Fermi operators and on leptoquarks using LEP 2 and lower energy data are given in Refs. 173 and 261. Constraints on various types of new physics are reviewed in Refs. [7,117,142,158,262,263], and implications for the LHC in Ref. 264.

An alternate formalism [265] defines parameters,  $\epsilon_1$ ,  $\epsilon_2$ ,  $\epsilon_3$ , and  $\epsilon_b$  in terms of the specific observables  $M_W/M_Z$ ,  $\Gamma_{\ell\ell}$ ,  $A_{FB}^{(0,\ell)}$ , and  $R_b$ . The definitions coincide with those for  $\tilde{c}_i$  in Eqs. (10.66) and (10.67) for physics which affects gauge self-energies only, but the  $\epsilon$ 's now parametrize arbitrary types of new physics. However, the  $\epsilon$ 's are not related to other observables unless additional model-dependent assumptions are made. Another approach [266] parametrizes new physics in terms of gauge-invariant sets of operators. It is especially powerful in studying the effects of new physics on non-Abelian gauge vertices. The most general approach introduces deviation vectors [262]. Each type of new physics defines a deviation vector, the components of which are the deviations of each observable from its SM prediction, normalized to the experimental uncertainty. The length (direction) of the vector represents the strength (type) of new physics.

For a particularly well motivated and explored type of physics beyond the SM, see the note on “The  $Z'$  Searches” in the Gauge & Higgs Boson Particle Listings.

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#### References:

1. S. L. Glashow, Nucl. Phys. **22**, 579 (1961);  
S. Weinberg, Phys. Rev. Lett. **19**, 1264 (1967);  
A. Salam, p. 367 of *Elementary Particle Theory*, ed.  
N. Svartholm (Almqvist and Wiksells, Stockholm, 1969);  
S.L. Glashow, J. Iliopoulos, and L. Maiani, Phys. Rev. **D2**,  
1285 (1970).
2. CKMfitter Group: J. Charles *et al.*, Eur. Phys. J. **C41**,  
1 (2005); updated results and plots are available at  
<http://ckmfinder.in2p3.fr>.
3. For reviews, see the Section on “Higgs Bosons: Theory and  
Searches” in this Review;  
J. Gunion *et al.*, *The Higgs Hunter's Guide*, (Addison-Wesley,  
Redwood City, 1990);  
M. Carena and H.E. Haber, Prog. in Part. Nucl. Phys. **50**, 63  
(2003);  
A. Djouadi, Phys. Reports **457**, 1 (2008).
4. For reviews, see E.D. Commins and P.H. Bucksbaum, *Weak  
Interactions of Leptons and Quarks*, (Cambridge Univ. Press,  
1983);  
G. Barbiellini and C. Santoni, Riv. Nuovo Cimento **9(2)**, 1  
(1986);  
N. Severijns, M. Beck, and O. Naviliat-Cuncic, Rev. Mod. Phys.  
**78**, 991 (2006);  
W. Fetscher and H.J. Gerber, p. 657 of Ref. 5;  
J. Deutsch and P. Quin, p. 706 of Ref. 5;  
P. Herczeg, p. 786 of Ref. 5;  
P. Herczeg, Prog. in Part. Nucl. Phys. **46**, 413 (2001).
5. *Precision Tests of the Standard Electroweak Model*, ed.  
P. Langacker (World Scientific, Singapore, 1995).
6. J. Erler and M.J. Ramsey-Musolf, Prog. in Part. Nucl. Phys.  
**54**, 351 (2005).
7. P. Langacker, *The Standard Model and Beyond*, (CRC Press,  
New York, 2009).
8. T. Kinoshita, *Quantum Electrodynamics*, (World Scientific,  
Singapore, 1990).
9. S. G. Karshenboim, Phys. Reports **422**, 1 (2005).

10. ALEPH, DELPHI, L3, OPAL, SLD, LEP Electroweak Working Group, SLD Electroweak and Heavy Flavour Groups: S. Schael *et al.*, Phys. Reports **427**, 257 (2006); for updates see <http://lepewwg.web.cern.ch/LEPEWWG/>.
11. MuLan: D.M. Webber *et al.*, Phys. Rev. Lett. **106**, 041803 (2011).
12. T. Kinoshita and A. Sirlin, Phys. Rev. **113**, 1652 (1959).
13. T. van Ritbergen and R.G. Stuart, Nucl. Phys. **B564**, 343 (2000);  
M. Steinhauser and T. Seidensticker, Phys. Lett. **B467**, 271 (1999).
14. Y. Nir, Phys. Lett. **B221**, 184 (1989).
15. A. Pak and A. Czarnecki, Phys. Rev. Lett. **100**, 241807 (2008).
16. P.J. Mohr, B.N. Taylor, and D.B. Newell, [arXiv:1507.07956](https://arxiv.org/abs/1507.07956) [physics.atom-ph].
17. For a review, see S. Mele, [arXiv:hep-ex/0610037](https://arxiv.org/abs/hep-ex/0610037).
18. S. Fanchiotti, B. Kniehl, and A. Sirlin, Phys. Rev. **D48**, 307 (1993) and references therein.
19. J. Erler, Phys. Rev. **D59**, 054008 (1999).
20. F. Jegerlehner, [arXiv:1511.04473](https://arxiv.org/abs/1511.04473) [hep-ph]; the quoted value uses additional information thankfully provided to us by the author in a private communication.
21. J. Erler, [hep-ph/0005084](https://arxiv.org/abs/hep-ph/0005084).
22. M. Steinhauser, Phys. Lett. **B429**, 158 (1998).
23. K.G. Chetyrkin, J.H. Kühn, and M. Steinhauser, Nucl. Phys. **B482**, 213 (1996).
24. B.V. Geshkenbein and V.L. Morgunov, Phys. Lett. **B352**, 456 (1995).
25. M.L. Swartz, Phys. Rev. **D53**, 5268 (1996).
26. N.V. Krasnikov and R. Rodenberg, Nuovo Cimento **111A**, 217 (1998).
27. J.H. Kühn and M. Steinhauser, Phys. Lett. **B437**, 425 (1998).
28. S. Groote *et al.*, Phys. Lett. **B440**, 375 (1998).
29. A.D. Martin, J. Outhwaite, and M.G. Ryskin, Phys. Lett. **B492**, 69 (2000).
30. J.F. de Troconiz and F.J. Yndurain, Phys. Rev. **D65**, 093002 (2002).
31. M. Davier *et al.*, Eur. Phys. J. **C71**, 1515 (2011).
32. H. Burkhardt and B. Pietrzyk, Phys. Rev. **D84**, 037502 (2011).
33. K. Hagiwara *et al.*, J. Phys. **G38**, 085003 (2011).
34. OPAL: K. Ackerstaff *et al.*, Eur. Phys. J. **C7**, 571 (1999).
35. CLEO: S. Anderson *et al.*, Phys. Rev. **D61**, 112002 (2000).
36. ALEPH: S. Schael *et al.*, Phys. Reports **421**, 191 (2005).
37. Belle: M. Fujikawa *et al.*, Phys. Rev. **D78**, 072006 (2008).
38. CMD-2: R.R. Akhmetshin *et al.*, Phys. Lett. **B578**, 285 (2004);  
CMD-2: V.M. Aulchenko *et al.*, JETP Lett. **82**, 743 (2005);  
CMD-2: R.R. Akhmetshin *et al.*, JETP Lett. **84**, 413 (2006);  
CMD-2: R.R. Akhmetshin *et al.*, Phys. Lett. **B648**, 28 (2007).
39. SND: M.N. Achasov *et al.*, Sov. Phys. JETP **103**, 380 (2006).
40. A.B. Arbuzov *et al.*, JHEP **9812**, 009 (1998);  
S. Binner, J.H. Kühn, and K. Melnikov, Phys. Lett. **B459**, 279 (1999).
41. BaBar: B. Aubert *et al.*, Phys. Rev. **D70**, 072004 (2004);  
BaBar: B. Aubert *et al.*, Phys. Rev. **D71**, 052001 (2005);  
BaBar: B. Aubert *et al.*, Phys. Rev. **D73**, 052003 (2006);  
BaBar: B. Aubert *et al.*, Phys. Rev. **D76**, 092005 (2007);  
BaBar: B. Aubert *et al.*, Phys. Rev. Lett. **103**, 231801 (2009).
42. KLOE: F. Ambrosino *et al.*, Phys. Lett. **B670**, 285 (2009);  
KLOE: F. Ambrosino *et al.*, Phys. Lett. **B700**, 102 (2011).
43. See *e.g.*, CMD and OLYA: L.M. Barkov *et al.*, Nucl. Phys. **B256**, 365 (1985).
44. V.A. Novikov *et al.*, Phys. Reports **41**, 1 (1978).
45. J. Erler and M. Luo, Phys. Lett. **B558**, 125 (2003).
46. K.G. Chetyrkin *et al.*, Theor. Math. Phys. **170**, 217 (2012).
47. B. Dehnadi *et al.*, JHEP **1309**, 103 (2013).
48. Tevatron Electroweak Working Group for the CDF and DØ Collaborations: [arXiv:1407.2682](https://arxiv.org/abs/1407.2682) [hep-ex].
49. Tevatron Electroweak and Top Physics LHC Working Groups for the ATLAS, CDF, CMS and DØ Collaborations: [arXiv:1403.4427](https://arxiv.org/abs/1403.4427) [hep-ex].
50. M. Beneke, Phys. Reports **317**, 1 (1999).
51. K.G. Chetyrkin and M. Steinhauser, Nucl. Phys. **B573**, 617 (2000);  
K. Melnikov and T. van Ritbergen, Phys. Lett. **B482**, 99 (2000).
52. P. Marquard *et al.*, Phys. Rev. Lett. **114**, 142002 (2015).
53. DØ: V. M. Abazov *et al.*, Phys. Rev. Lett. **113**, 032002 (2014).
54. CMS: <http://cds.cern.ch/record/1690093/files/TOP-14-001-pas.pdf>.
55. For a recent review, see J. Erler, [arXiv:1412.4435](https://arxiv.org/abs/1412.4435) [hep-ph].
56. A. Sirlin, Phys. Rev. **D22**, 971 (1980);  
D.C. Kennedy *et al.*, Nucl. Phys. **B321**, 83 (1989);  
D.Yu. Bardin *et al.*, Z. Phys. **C44**, 493 (1989);  
W. Hollik, Fortsch. Phys. **38**, 165 (1990);  
for reviews, see W. Hollik, pp. 37 and 117, and W. Marciano, p. 170 of Ref. 5.
57. W.J. Marciano and J.L. Rosner, Phys. Rev. Lett. **65**, 2963 (1990).
58. G. Degrassi, S. Fanchiotti, and A. Sirlin, Nucl. Phys. **B351**, 49 (1991).
59. K.S. Kumar *et al.*, Ann. Rev. Nucl. and Part. Sci. **63**, 237 (2013).
60. G. Degrassi and A. Sirlin, Nucl. Phys. **B352**, 342 (1991).
61. P. Gambino and A. Sirlin, Phys. Rev. **D49**, 1160 (1994).
62. ZFITTER: A.B. Arbuzov *et al.*, Comput. Phys. Commun. **174**, 728 (2006) and references therein.
63. R. Barbieri *et al.*, Nucl. Phys. **B409**, 105 (1993);  
J. Fleischer, O.V. Tarasov, and F. Jegerlehner, Phys. Lett. **B319**, 249 (1993).
64. G. Degrassi, P. Gambino, and A. Vicini, Phys. Lett. **B383**, 219 (1996);  
G. Degrassi, P. Gambino, and A. Sirlin, Phys. Lett. **B394**, 188 (1997).
65. A. Freitas *et al.*, Phys. Lett. **B495**, 338 (2000) and *ibid.* **570**, 260(E) (2003);  
M. Awramik and M. Czakon, Phys. Lett. **B568**, 48 (2003).
66. A. Freitas *et al.*, Nucl. Phys. **B632**, 189 (2002) and *ibid.* **666**, 305(E) (2003);  
M. Awramik and M. Czakon, Phys. Rev. Lett. **89**, 241801 (2002);  
A. Onishchenko and O. Veretin, Phys. Lett. **B551**, 111 (2003).
67. M. Awramik *et al.*, Phys. Rev. Lett. **93**, 201805 (2004);  
W. Hollik, U. Meier, and S. Uccirati, Nucl. Phys. **B731**, 213 (2005).
68. M. Awramik, M. Czakon, and A. Freitas, Phys. Lett. **B642**, 563 (2006);  
W. Hollik, U. Meier, and S. Uccirati, Nucl. Phys. **B765**, 154 (2007).
69. G. Degrassi, P. Gambino and P.P. Giardino, JHEP **1505**, 154 (2015).
70. M. Awramik *et al.*, Phys. Rev. **D69**, 053006 (2004).
71. A. Djouadi and C. Verzegnassi, Phys. Lett. **B195**, 265 (1987);  
A. Djouadi, Nuovo Cimento **100A**, 357 (1988).
72. K.G. Chetyrkin, J.H. Kühn, and M. Steinhauser, Phys. Lett. **B351**, 331 (1995).
73. Y. Schröder and M. Steinhauser, Phys. Lett. **B622**, 124 (2005);  
K.G. Chetyrkin *et al.*, Phys. Rev. Lett. **97**, 102003 (2006);  
R. Boughezal and M. Czakon, Nucl. Phys. **B755**, 221 (2006);  
L. Avdeev *et al.*, Phys. Lett. **B336**, 560 (1994) and *ibid.* **B349**, 597(E) (1995).
74. B.A. Kniehl, J.H. Kühn, and R.G. Stuart, Phys. Lett. **B214**, 621 (1988);  
B.A. Kniehl, Nucl. Phys. **B347**, 86 (1990);  
F. Halzen and B.A. Kniehl, Nucl. Phys. **B353**, 567 (1991);  
A. Djouadi and P. Gambino, Phys. Rev. **D49**, 3499 (1994) and *ibid.* **53**, 4111(E) (1996).
75. A. Anselm, N. Dombey, and E. Leader, Phys. Lett. **B312**, 232 (1993).
76. K.G. Chetyrkin, J.H. Kühn, and M. Steinhauser, Phys. Rev. Lett. **75**, 3394 (1995).
77. J.J. van der Bij *et al.*, Phys. Lett. **B498**, 156 (2001);  
M. Faisst *et al.*, Nucl. Phys. **B665**, 649 (2003).



78. R. Boughezal, J.B. Tausk, and J.J. van der Bij, Nucl. Phys. **B725**, 3 (2005).
79. M. Awramik, M. Czakon, and A. Freitas, JHEP **0611**, 048 (2006).
80. J. Erler and S. Su, Prog. in Part. Nucl. Phys. **71**, 119 (2013).
81. J.M. Conrad, M.H. Shaevitz, and T. Bolton, Rev. Mod. Phys. **70**, 1341 (1998).
82. Z. Berezhiani and A. Rossi, Phys. Lett. **B535**, 207 (2002); S. Davidson *et al.*, JHEP **0303**, 011 (2003); A. Friedland, C. Lunardini and C. Pena-Garay, Phys. Lett. **B594**, 347 (2004).
83. J. Panman, p. 504 of Ref. 5.
84. CHARM: J. Dorenbosch *et al.*, Z. Phys. **C41**, 567 (1989).
85. CALO: L.A. Ahrens *et al.*, Phys. Rev. **D41**, 3297 (1990).
86. CHARM II: P. Vilain *et al.*, Phys. Lett. **B335**, 246 (1994).
87. ILM: R.C. Allen *et al.*, Phys. Rev. **D47**, 11 (1993); LSND: L.B. Auerbach *et al.*, Phys. Rev. **D63**, 112001 (2001).
88. TEXONO: M. Deniz *et al.*, Phys. Rev. **D81**, 072001 (2010).
89. For reviews, see G.L. Fogli and D. Haidt, Z. Phys. **C40**, 379 (1988); F. Perrier, p. 385 of Ref. 5.
90. CDHS: A. Blondel *et al.*, Z. Phys. **C45**, 361 (1990).
91. CHARM: J.V. Allaby *et al.*, Z. Phys. **C36**, 611 (1987).
92. CCFR: K.S. McFarland *et al.*, Eur. Phys. J. **C1**, 509 (1998).
93. R.M. Barnett, Phys. Rev. **D14**, 70 (1976); H. Georgi and H.D. Politzer, Phys. Rev. **D14**, 1829 (1976).
94. LAB-E: S.A. Rabinowitz *et al.*, Phys. Rev. Lett. **70**, 134 (1993).
95. E.A. Paschos and L. Wolfenstein, Phys. Rev. **D7**, 91 (1973).
96. NuTeV: G.P. Zeller *et al.*, Phys. Rev. Lett. **88**, 091802 (2002).
97. D. Mason *et al.*, Phys. Rev. Lett. **99**, 192001 (2007).
98. S. Kretzer, D. Mason, and F. Olness, Phys. Rev. **D65**, 074010 (2002).
99. J. Pumplin *et al.*, JHEP **0207**, 012 (2002); S. Kretzer *et al.*, Phys. Rev. Lett. **93**, 041802 (2004).
100. E. Sather, Phys. Lett. **B274**, 433 (1992); E.N. Rodionov, A.W. Thomas, and J.T. Londergan, Mod. Phys. Lett. **A9**, 1799 (1994).
101. A.D. Martin *et al.*, Eur. Phys. J. **C35**, 325 (2004).
102. J.T. Londergan and A.W. Thomas, Phys. Rev. **D67**, 111901 (2003).
103. W. Bentz *et al.*, Phys. Lett. **B693**, 462 (2010).
104. M. Glück, P. Jimenez-Delgado, and E. Reya, Phys. Rev. Lett. **95**, 022002 (2005).
105. S. Kumano, Phys. Rev. **D66**, 111301 (2002); S.A. Kulagin, Phys. Rev. **D67**, 091301 (2003); S.J. Brodsky, I. Schmidt, and J.J. Yang, Phys. Rev. **D70**, 116003 (2004); M. Hirai, S. Kumano, and T. H. Nagai, Phys. Rev. **D71**, 113007 (2005); G.A. Miller and A.W. Thomas, Int. J. Mod. Phys. A **20**, 95 (2005).
106. I.C. Cloet, W. Bentz, and A.W. Thomas, Phys. Rev. Lett. **102**, 252301 (2009).
107. K.P.O. Diener, S. Dittmaier, and W. Hollik, Phys. Rev. **D69**, 073005 (2004); A.B. Arbuzov, D.Y. Bardin, and L.V. Kalinovskaya, JHEP **0506**, 078 (2005); K. Park, U. Baur, and D. Wackeroth, arXiv:0910.5013 [hep-ph].
108. K.P.O. Diener, S. Dittmaier, and W. Hollik, Phys. Rev. **D72**, 093002 (2005).
109. B.A. Dobrescu and R.K. Ellis, Phys. Rev. **D69**, 114014 (2004).
110. For a review, see S. Davidson *et al.*, JHEP **0202**, 037 (2002).
111. J. Erler *et al.*, Ann. Rev. Nucl. and Part. Sci. **64**, 269 (2014).
112. SSF: C.Y. Prescott *et al.*, Phys. Lett. **B84**, 524 (1979).
113. JLab-PV DIS: D. Wang *et al.*, Nature **506**, 67 (2014) and Phys. Rev. **C91**, 045506 (2015).
114. E.J. Beise, M.L. Pitt, and D.T. Spayde, Prog. in Part. Nucl. Phys. **54**, 289 (2005).
115. S.L. Zhu *et al.*, Phys. Rev. **D62**, 033008 (2000).
116. P. Souder, p. 599 of Ref. 5.
117. P. Langacker, p. 883 of Ref. 5.
118. R.D. Young *et al.*, Phys. Rev. Lett. **99**, 122003 (2007).
119. D.S. Armstrong and R.D. McKeown, Ann. Rev. Nucl. and Part. Sci. **62**, 337 (2012).
120. E. Derman and W.J. Marciano, Annals Phys. **121**, 147 (1979).
121. E158: P.L. Anthony *et al.*, Phys. Rev. Lett. **95**, 081601 (2005).
122. J. Erler and M.J. Ramsey-Musolf, Phys. Rev. **D72**, 073003 (2005).
123. A. Czarnecki and W.J. Marciano, Int. J. Mod. Phys. A **15**, 2365 (2000).
124. C. Bouchiat and C.A. Piketty, Phys. Lett. **B128**, 73 (1983).
125. K.S. McFarland, in the *Proceedings of DIS* 2008.
126. Qweak: M.T. Gericke *et al.*, AIP Conf. Proc. **1149**, 237 (2009).
127. Qweak: D. Androic *et al.*, Phys. Rev. Lett. **111**, 141803 (2013); the implications are discussed in Ref. 142.
128. For reviews and references to earlier work, see M.A. Bouchiat and L. Pottier, Science **234**, 1203 (1986); B.P. Masterson and C.E. Wieman, p. 545 of Ref. 5.
129. Cesium (Boulder): C.S. Wood *et al.*, Science **275**, 1759 (1997).
130. Cesium (Paris): J. Guéna, M. Lintz, and M.A. Bouchiat, Phys. Rev. **A71**, 042108 (2005).
131. Thallium (Oxford): N.H. Edwards *et al.*, Phys. Rev. Lett. **74**, 2654 (1995); Thallium (Seattle): P.A. Vetter *et al.*, Phys. Rev. Lett. **74**, 2658 (1995).
132. Lead (Seattle): D.M. Meekhof *et al.*, Phys. Rev. Lett. **71**, 3442 (1993).
133. Bismuth (Oxford): M.J.D. MacPherson *et al.*, Phys. Rev. Lett. **67**, 2784 (1991).
134. P.G. Blunden, W. Melnitchouk and A.W. Thomas, Phys. Rev. Lett. **109**, 262301 (2012).
135. V.A. Dzuba, V.V. Flambaum, and O.P. Sushkov, Phys. Lett. **141A**, 147 (1989); S.A. Blundell, J. Sapirstein, and W.R. Johnson, Phys. Rev. **D45**, 1602 (1992); For reviews, see S.A. Blundell, W.R. Johnson, and J. Sapirstein, p. 577 of Ref. 5; J.S.M. Ginges and V.V. Flambaum, Phys. Reports **397**, 63 (2004); J. Guéna, M. Lintz, and M. A. Bouchiat, Mod. Phys. Lett. **A20**, 375 (2005); A. Derevianko and S.G. Porsev, Eur. Phys. J. A **32**, 517 (2007).
136. V.A. Dzuba, V.V. Flambaum, and O.P. Sushkov, Phys. Rev. **A56**, R4357 (1997).
137. S.C. Bennett and C.E. Wieman, Phys. Rev. Lett. **82**, 2484 (1999).
138. M.A. Bouchiat and J. Guéna, J. Phys. (France) **49**, 2037 (1988).
139. S. G. Porsev, K. Beloy, and A. Derevianko, Phys. Rev. Lett. **102**, 181601 (2009).
140. V.A. Dzuba *et al.*, Phys. Rev. Lett. **109**, 203003 (2012); B. M. Roberts, V. A. Dzuba and V. V. Flambaum, Ann. Rev. Nucl. and Part. Sci. **65**, 63 (2015).
141. A. Derevianko, Phys. Rev. Lett. **85**, 1618 (2000); V.A. Dzuba, C. Harabati, and W.R. Johnson, Phys. Rev. **A63**, 044103 (2001); M.G. Kozlov, S.G. Porsev, and I.I. Tupitsyn, Phys. Rev. Lett. **86**, 3260 (2001); W.R. Johnson, I. Bednyakov, and G. Soff, Phys. Rev. Lett. **87**, 233001 (2001); A.I. Milstein and O.P. Sushkov, Phys. Rev. **A66**, 022108 (2002); V.A. Dzuba, V.V. Flambaum, and J.S. Ginges, Phys. Rev. **D66**, 076013 (2002); M.Y. Kuchiev and V.V. Flambaum, Phys. Rev. Lett. **89**, 283002 (2002); A.I. Milstein, O.P. Sushkov, and I.S. Terekhov, Phys. Rev. Lett. **89**, 283003 (2002); V.V. Flambaum and J.S.M. Ginges, Phys. Rev. **A72**, 052115 (2005).
142. J. Erler, A. Kurylov, and M.J. Ramsey-Musolf, Phys. Rev. **D68**, 016006 (2003).



143. V.A. Dzuba *et al.*, J. Phys. **B20**, 3297 (1987).
144. Ya.B. Zel'dovich, Sov. Phys. JETP **6**, 1184 (1958);  
V.V. Flambaum and D.W. Murray, Phys. Rev. **C56**, 1641 (1997);  
W.C. Haxton and C.E. Wieman, Ann. Rev. Nucl. Part. Sci. **51**, 261 (2001).
145. J.L. Rosner, Phys. Rev. **D53**, 2724 (1996).
146. S.J. Pollock, E.N. Fortson, and L. Wilets, Phys. Rev. **C46**, 2587 (1992);  
B.Q. Chen and P. Vogel, Phys. Rev. **C48**, 1392 (1993).
147. R.W. Dunford and R.J. Holt, J. Phys. **G34**, 2099 (2007).
148. O.O. Versolato *et al.*, Hyperfine Interact. **199**, 9 (2011).
149. W. Hollik and G. Duckeck, Springer Tracts Mod. Phys. **162**, 1 (2000).
150. D.Yu. Bardin *et al.*, hep-ph/9709229.
151. A. Czarnecki and J.H. Kühn, Phys. Rev. Lett. **77**, 3955 (1996);  
R. Harlander, T. Seidensticker, and M. Steinhauser, Phys. Lett. **B426**, 125 (1998);  
J. Fleischer *et al.*, Phys. Lett. **B459**, 625 (1999).
152. M. Awramik *et al.*, Nucl. Phys. **B813**, 174 (2009).
153. A. Freitas, Phys. Lett. **B730**, 50 (2014) and JHEP **1404**, 070 (2014).
154. B.W. Lynn and R.G. Stuart, Nucl. Phys. **B253**, 216 (1985).
155. *Physics at LEP*, ed. J. Ellis and R. Peccei, CERN 86-02, Vol. 1.
156. PETRA: S.L. Wu, Phys. Reports **107**, 59 (1984);  
C. Kiesling, *Tests of the Standard Theory of Electroweak Interactions*, (Springer-Verlag, New York, 1988);  
R. Marshall, Z. Phys. **C43**, 607 (1989);  
Y. Mori *et al.*, Phys. Lett. **B218**, 499 (1989);  
D. Haidt, p. 203 of Ref. 5.
157. For reviews, see D. Schaile, p. 215, and A. Blondel, p. 277 of Ref. 5; P. Langacker [7]; and S. Riemann [158].
158. S. Riemann, Rept. on Prog. in Phys. **73**, 126201 (2010).
159. SLD: K. Abe *et al.*, Phys. Rev. Lett. **84**, 5945 (2000).
160. SLD: K. Abe *et al.*, Phys. Rev. Lett. **85**, 5059 (2000).
161. SLD: K. Abe *et al.*, Phys. Rev. Lett. **86**, 1162 (2001).
162. SLD: K. Abe *et al.*, Phys. Rev. Lett. **78**, 17 (1997).
163. P.A. Grassi, B.A. Kniehl, and A. Sirlin, Phys. Rev. Lett. **86**, 389 (2001).
164. DØ: V. M. Abazov *et al.*, Phys. Rev. **D115**, 041801 (2015).
165. CDF: T. A. Aaltonen *et al.*, Phys. Rev. **D89**, 072005 (2014).
166. CDF: <http://www-cdf.fnal.gov/physics/ewk/2015/zAfb9ee/pnAfb9ee.pdf>.
167. DØ: V.M. Abazov *et al.*, Phys. Rev. **D84**, 012007 (2011).
168. CDF: D. Acosta *et al.*, Phys. Rev. **D71**, 052002 (2005).
169. H1: A. Aktas *et al.*, Phys. Lett. **B632**, 35 (2006);  
H1 and ZEUS: Z. Zhang, Nucl. Phys. Proc. Suppl. **191**, 271 (2009).
170. ATLAS: G. Aad *et al.*, JHEP **1509**, 049 (2015).
171. CMS: S. Chatrchyan *et al.*, Phys. Rev. **D84**, 112002 (2011).
172. LHCb: R. Aaij *et al.*, arXiv:1509.07645 [hep-ex].
173. ALEPH, DELPHI, L3, OPAL, and LEP Electroweak Working Group: S. Schael *et al.*, Phys. Reports **532**, 119 (2013).
174. ALEPH, DELPHI, L3, OPAL, and LEP Electroweak Working Group: J. Alcaraz *et al.*, hep-ex/0612034.
175. A. Leike, T. Riemann, and J. Rose, Phys. Lett. **B273**, 513 (1991);  
T. Riemann, Phys. Lett. **B293**, 451 (1992).
176. A comprehensive report and further references can be found in K.G. Chetyrkin, J.H. Kühn, and A. Kwiatkowski, Phys. Reports **277**, 189 (1996).
177. J. Schwinger, *Particles, Sources, and Fields*, Vol. II, (Addison-Wesley, New York, 1973);  
K.G. Chetyrkin, A.L. Kataev, and F.V. Tkachev, Phys. Lett. **B85**, 277 (1979);  
M. Dine and J. Sapiirstein, Phys. Rev. Lett. **43**, 668 (1979);  
W. Celmaster and R.J. Gonsalves, Phys. Rev. Lett. **44**, 560 (1980);  
S.G. Gorishnii, A.L. Kataev, and S.A. Larin, Phys. Lett. **B259**, 144 (1991);  
L.R. Surguladze, M.A. Samuel, Phys. Rev. Lett. **66**, 560 (1991) and *ibid.* 2416(E).
178. P.A. Baikov, K.G. Chetyrkin, and J.H. Kühn, Phys. Rev. Lett. **101**, 012002 (2008).
179. B.A. Kniehl and J.H. Kühn, Nucl. Phys. **B329**, 547 (1990);  
K.G. Chetyrkin and A. Kwiatkowski, Phys. Lett. **B319**, 307 (1993);  
S.A. Larin, T. van Ritbergen, and J.A.M. Vermaseren, Phys. Lett. **B320**, 159 (1994);  
K.G. Chetyrkin and O.V. Tarasov, Phys. Lett. **B327**, 114 (1994).
180. A.L. Kataev, Phys. Lett. **B287**, 209 (1992).
181. D. Albert *et al.*, Nucl. Phys. **B166**, 460 (1980);  
F. Jegerlehner, Z. Phys. **C32**, 425 (1986);  
A. Djouadi, J.H. Kühn, and P.M. Zerwas, Z. Phys. **C46**, 411 (1990);  
A. Borrelli *et al.*, Nucl. Phys. **B333**, 357 (1990).
182. A.A. Akhundov, D.Yu. Bardin, and T. Riemann, Nucl. Phys. **B276**, 1 (1986);  
W. Beenakker and W. Hollik, Z. Phys. **C40**, 141 (1988);  
B.W. Lynn and R.G. Stuart, Phys. Lett. **B352**, 676 (1990);  
J. Bernabeu, A. Pich, and A. Santamaria, Phys. Lett. **B200**, 569 (1988) and Nucl. Phys. **B363**, 326 (1991).
183. ATLAS: G. Aad *et al.*, Phys. Lett. **B716**, 1 (2012);  
CMS: S. Chatrchyan *et al.*, Phys. Lett. **B716**, 30 (2012);  
CMS: S. Chatrchyan *et al.*, JHEP **1306**, 081 (2013).
184. ATLAS and CMS: G. Aad *et al.*, Phys. Rev. Lett. **114**, 191803 (2015) and [cds.cern.ch/record/2052552/files/ATLAS-CONF-2015-044.pdf](https://cds.cern.ch/record/2052552/files/ATLAS-CONF-2015-044.pdf).
185. A. Djouadi, J. Kalinowski, and M. Spira, Comp. Phys. Comm. **108**, 56 (1998).
186. H. Davoudiasl, H.-S. Lee, and W.J. Marciano, Phys. Rev. **D86**, 095009 (2012).
187. E. Braaten, S. Narison, and A. Pich, Nucl. Phys. **B373**, 581 (1992).
188. M. Davier *et al.*, Eur. Phys. J. **C74**, 2803 (2014).
189. F. Le Diberder and A. Pich, Phys. Lett. **B286**, 147 (1992).
190. M. Beneke and M. Jamin, JHEP **0809**, 044 (2008).
191. E. Braaten and C.S. Li, Phys. Rev. **D42**, 3888 (1990).
192. J. Erler, Rev. Mex. Fis. **50**, 200 (2004).
193. D. Boito *et al.*, Phys. Rev. **D85**, 093015 (2012).
194. D. Boito *et al.*, Phys. Rev. **D91**, 034003 (2015).
195. J. Erler, arXiv:1102.5520 [hep-ph].
196. E821: G.W. Bennett *et al.*, Phys. Rev. Lett. **92**, 161802 (2004).
197. T. Aoyama *et al.*, Phys. Rev. Lett. **109**, 111808 (2012);  
T. Aoyama *et al.*, PTEP **2012**, 01A107 (2012);  
P. Baikov, A. Maier and P. Marquard, Nucl. Phys. **B877**, 647 (2013).
198. G. Li, R. Mendel, and M.A. Samuel, Phys. Rev. **D47**, 1723 (1993);  
S. Laporta and E. Remiddi, Phys. Lett. **B301**, 440 (1993);  
S. Laporta and E. Remiddi, Phys. Lett. **B379**, 283 (1996);  
A. Czarnecki and M. Skrzypek, Phys. Lett. **B449**, 354 (1999).
199. J. Erler and M. Luo, Phys. Rev. Lett. **87**, 071804 (2001).
200. S.J. Brodsky and J.D. Sullivan, Phys. Rev. **D156**, 1644 (1967);  
T. Burnett and M.J. Levine, Phys. Lett. **B24**, 467 (1967);  
R. Jackiw and S. Weinberg, Phys. Rev. **D5**, 2473 (1972);  
I. Bars and M. Yoshimura, Phys. Rev. **D6**, 374 (1972);  
K. Fujikawa, B.W. Lee, and A.I. Sanda, Phys. Rev. **D6**, 2923 (1972);  
G. Altarelli, N. Cabibbo, and L. Maiani, Phys. Lett. **B40**, 415 (1972);  
W.A. Bardeen, R. Gastmans, and B.E. Laurup, Nucl. Phys. **B46**, 315 (1972).
201. T.V. Kukhto *et al.*, Nucl. Phys. **B371**, 567 (1992);  
S. Peris, M. Perrottet, and E. de Rafael, Phys. Lett. **B355**, 523 (1995);  
A. Czarnecki, B. Krause, and W.J. Marciano, Phys. Rev. **D52**, 2619 (1995);  
A. Czarnecki, B. Krause, and W.J. Marciano, Phys. Rev. Lett. **76**, 3267 (1996);

- S. Heinemeyer, D. Stöckinger, and G. Weiglein, Nucl. Phys. **B699**, 103 (2004).
202. G. Degrandi and G. Giudice, Phys. Rev. **D58**, 053007 (1998); A. Czarnecki, W.J. Marciano and A. Vainshtein, Phys. Rev. **D67**, 073006 (2003) and *ibid.* **D73**, 119901(E) (2006).
203. For reviews, see M. Davier and W.J. Marciano, Ann. Rev. Nucl. Part. Sci. **54**, 115 (2004); J.P. Miller, E. de Rafael, and B.L. Roberts, Rept. Prog. Phys. **70**, 795 (2007); F. Jegerlehner and A. Nyffeler, Phys. Reports **477**, 1 (2009).
204. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **61**, 1815 (1988).
205. F. Jegerlehner and R. Szafron, Eur. Phys. J. **C71**, 1632 (2011).
206. F. Burger *et al.*, PoS LATTICE **2013**, 301 (2014).
207. K. Melnikov and A. Vainshtein, Phys. Rev. **D70**, 113006 (2004).
208. J. Erler and G. Toledo Sánchez, Phys. Rev. Lett. **97**, 161801 (2006).
209. A. Nyffeler, Phys. Rev. **D79**, 073012 (2009).
210. J. Prades, E. de Rafael, and A. Vainshtein, Adv. Ser. Direct. High Energy Phys. **20**, 303 (2009).
211. M. Knecht and A. Nyffeler, Phys. Rev. **D65**, 073034 (2002).
212. M. Hayakawa and T. Kinoshita, hep-ph/0112102; J. Bijnens, E. Pallante, and J. Prades, Nucl. Phys. **B626**, 410 (2002); A recent discussion is in J. Bijnens and J. Prades, Mod. Phys. Lett. **A22**, 767 (2007).
213. T. Blum *et al.*, arXiv:1510.07100 [hep-lat].
214. B. Krause, Phys. Lett. **B390**, 392 (1997).
215. A. Kurz, T. Liu, P. Marquard and M. Steinhauser, Phys. Lett. **B734**, 144 (2014).
216. G. Colangelo *et al.*, Phys. Lett. **B735**, 90 (2014).
217. J.L. Lopez, D.V. Nanopoulos, and X. Wang, Phys. Rev. **D49**, 366 (1994); for recent reviews, see Ref. 203.
218. U. Amaldi *et al.*, Phys. Rev. **D36**, 1385 (1987); G. Costa *et al.*, Nucl. Phys. **B297**, 244 (1988); P. Langacker and M. Luo, Phys. Rev. **D44**, 817 (1991); J. Erler and P. Langacker, Phys. Rev. **D52**, 441 (1995).
219. CDF and DØ: T. Aaltonen *et al.*, Phys. Rev. **D88**, 052018 (2013).
220. Tevatron Electroweak Working Group, CDF and DØ: arXiv:1003.2826 [hep-ex].
221. F. James and M. Roos, Comput. Phys. Commun. **10**, 343 (1975).
222. CMS: S. Chatrchyan *et al.*, Phys. Lett. **B728**, 496 (2014) and *ibid.* 526(E) (2014).
223. DØ: V.M. Abazov *et al.*, Phys. Rev. **D80**, 111107 (2009).
224. CMS: <http://cds.cern.ch/record/1690093/files/TOP-14-015-pas.pdf>.
225. P. Langacker and N. Polonsky, Phys. Rev. **D52**, 3081 (1995); J. Bagger, K.T. Matchev, and D. Pierce, Phys. Lett. **B348**, 443 (1995).
226. M. Veltman, Nucl. Phys. **B123**, 89 (1977); M. Chanowitz, M.A. Furman, and I. Hinchliffe, Phys. Lett. **B78**, 285 (1978); The two-loop correction has been obtained by J.J. van der Bij and F. Hoogeveen, Nucl. Phys. **B283**, 477 (1987).
227. P. Langacker and M. Luo, Phys. Rev. **D45**, 278 (1992) and refs. therein.
228. A. Denner, R.J. Guth, and J.H. Kühn, Phys. Lett. **B240**, 438 (1990); W. Grimus *et al.*, J. Phys. G **35**, 075001 (2008); H. E. Haber and D. O’Neil, Phys. Rev. **D83**, 055017 (2011).
229. S. Bertolini and A. Sirlin, Phys. Lett. **B257**, 179 (1991).
230. M. Peskin and T. Takeuchi, Phys. Rev. Lett. **65**, 964 (1990); M. Peskin and T. Takeuchi, Phys. Rev. **D46**, 381 (1992); M. Golden and L. Randall, Nucl. Phys. **B361**, 3 (1991).
231. D. Kennedy and P. Langacker, Phys. Rev. **D44**, 1591 (1991).
232. G. Altarelli and R. Barbieri, Phys. Lett. **B253**, 161 (1991).
233. B. Holdom and J. Terning, Phys. Lett. **B247**, 88 (1990).
234. B.W. Lynn, M.E. Peskin, and R.G. Stuart, p. 90 of Ref. 155.
235. An alternative formulation is given by K. Hagiwara *et al.*, Z. Phys. **C64**, 559 (1994), and *ibid.* **68**, 352(E) (1995); K. Hagiwara, D. Haidt, and S. Matsumoto, Eur. Phys. J. **C2**, 95 (1998).
236. I. Maksymyk, C.P. Burgess, and D. London, Phys. Rev. **D50**, 529 (1994); C.P. Burgess *et al.*, Phys. Lett. **B326**, 276 (1994); R. Barbieri, M. Frigeni, and F. Caravaglios, Phys. Lett. **B279**, 169 (1992).
237. R. Barbieri *et al.*, Nucl. Phys. **B703**, 127 (2004).
238. M.S. Carena *et al.*, Nucl. Phys. **B759**, 202 (2006).
239. K. Lane, hep-ph/0202255.
240. E. Gates and J. Terning, Phys. Rev. Lett. **67**, 1840 (1991); R. Sundrum and S.D.H. Hsu, Nucl. Phys. **B391**, 127 (1993); R. Sundrum, Nucl. Phys. **B395**, 60 (1993); M. Luty and R. Sundrum, Phys. Rev. Lett. **70**, 529 (1993); T. Appelquist and J. Terning, Phys. Lett. **B315**, 139 (1993); D.D. Dietrich, F. Sannino, and K. Tuominen, Phys. Rev. **D72**, 055001 (2005); N.D. Christensen and R. Shrock, Phys. Lett. **B632**, 92 (2006); M. Harada, M. Kurachi, and K. Yamawaki, Prog. Theor. Phys. **115**, 765 (2006).
241. H. Georgi, Nucl. Phys. **B363**, 301 (1991); M.J. Dugan and L. Randall, Phys. Lett. **B264**, 154 (1991).
242. R. Barbieri *et al.*, Nucl. Phys. **B341**, 309 (1990).
243. J. Erler and D.M. Pierce, Nucl. Phys. **B526**, 53 (1998); G.C. Cho and K. Hagiwara, Nucl. Phys. **B574**, 623 (2000); G. Altarelli *et al.*, JHEP **0106**, 018 (2001); S. Heinemeyer, W. Hollik, and G. Weiglein, Phys. Reports **425**, 265 (2006); S.P. Martin, K. Tobe, and J.D. Wells, Phys. Rev. **D71**, 073014 (2005); G. Marandella, C. Schappacher, and A. Strumia, Nucl. Phys. **B715**, 173 (2005); M.J. Ramsey-Musolf and S. Su, Phys. Reports **456**, 1 (2008); A. Djouadi, Phys. Reports **459**, 1 (2008); S. Heinemeyer *et al.*, JHEP **0804**, 039 (2008); O. Buchmueller *et al.*, Eur. Phys. J. **C72**, 2020 (2012).
244. J. Erler and P. Langacker, Phys. Rev. Lett. **105**, 031801 (2010).
245. CMS: S. Chatrchyan *et al.*, Phys. Rev. **D86**, 112003 (2012).
246. O. Eberhardt *et al.*, Phys. Rev. Lett. **109**, 241802 (2012); A. Djouadi and A. Lenz, Phys. Lett. **B715**, 310 (2012).
247. J. Maalampi and M. Roos, Phys. Reports **186**, 53 (1990).
248. For reviews, see the Section on “Grand Unified Theories” in this Review; P. Langacker, Phys. Reports **72**, 185 (1981); J.L. Hewett and T.G. Rizzo, Phys. Reports **183**, 193 (1989); J. Kang, P. Langacker and B.D. Nelson, Phys. Rev. **D77**, 035003 (2008); S.P. Martin, Phys. Rev. **D81**, 035004 (2010); P.W. Graham *et al.*, Phys. Rev. **D81**, 055016 (2010).
249. J. Erler, J.L. Feng, and N. Polonsky, Phys. Rev. Lett. **78**, 3063 (1997).
250. D. Choudhury, T.M.P. Tait, and C.E.M. Wagner, Phys. Rev. **D65**, 053002 (2002).
251. J. Erler and P. Langacker, Phys. Rev. Lett. **84**, 212 (2000).
252. P. Langacker and M. Plümacher, Phys. Rev. **D62**, 013006 (2000).
253. DELPHI: P. Abreu *et al.*, Eur. Phys. J. **C10**, 415 (1999).
254. P. Langacker and D. London, Phys. Rev. **D38**, 886 (1988); D. London, p. 951 of Ref. 5; a recent analysis is F. del Aguila, J. de Blas and M. Perez-Victoria, Phys. Rev. **D78**, 013010 (2008).
255. M. Chemtob, Prog. in Part. Nucl. Phys. **54**, 71 (2005); R. Barbier *et al.*, Phys. Reports **420**, 1 (2005).
256. C.T. Hill and E.H. Simmons, Phys. Reports **381**, 235 (2003); R. Contino, arXiv:1005.4269 [hep-ph]; A. Pich, I. Rosell and J. J. Sanz-Cillero, JHEP **1401**, 157 (2014).
257. K. Agashe *et al.*, JHEP **0308**, 050 (2003); M. Carena *et al.*, Phys. Rev. **D68**, 035010 (2003);

- I. Gogoladze and C. Macesanu, Phys. Rev. **D74**, 093012 (2006);  
I. Antoniadis, [hep-th/0102202](#);  
see also the note on “Extra Dimensions” in the Searches Particle Listings.
258. T. Han, H.E. Logan, and L.T. Wang, JHEP **0601**, 099 (2006);  
M. Perelstein, Prog. in Part. Nucl. Phys. **58**, 247 (2007).
259. E. Accomando *et al.*, [arXiv:hep-ph/0608079](#);  
V. Barger *et al.*, Phys. Rev. **D77**, 035005 (2008);  
W. Grimus *et al.*, Nucl. Phys. **B801**, 81 (2008);  
M.C. Chen, S. Dawson, and C.B. Jackson, Phys. Rev. **D78**, 093001 (2008);  
M. Maniatis, Int. J. Mod. Phys. **A25**, 3505 (2010);  
U. Ellwanger, C. Hugonie, and A.M. Teixeira, Phys. Reports **496**, 1 (2010);  
S. Kanemura and K. Yagyu, Phys. Rev. **D85**, 115009 (2012).
260. A. Barroso *et al.*, [arXiv:1304.5225 \[hep-ph\]](#);  
B. Grinstein and P. Uttayarat, JHEP **1306**, 094 (2013) and *ibid.* **1309**, 110(E) (2013);  
O. Eberhardt, U. Nierste and M. Wiebusch, JHEP **1307**, 118 (2013);  
A. Broggio *et al.*, JHEP **1411**, 058 (2014).
261. G.C. Cho, K. Hagiwara, and S. Matsumoto, Eur. Phys. J. **C5**, 155 (1998);  
K. Cheung, Phys. Lett. **B517**, 167 (2001);  
Z. Han and W. Skiba, Phys. Rev. **D71**, 075009 (2005).
262. P. Langacker, M. Luo, and A.K. Mann, Rev. Mod. Phys. **64**, 87 (1992);  
M. Luo, p. 977 of Ref. 5.
263. F.S. Merritt *et al.*, p. 19 of *Particle Physics: Perspectives and Opportunities: Report of the DPF Committee on Long Term Planning*, ed. R. Peccei *et al.* (World Scientific, Singapore, 1995).
264. D.E. Morrissey, T. Plehn, and T.M.P. Tait, Phys. Reports **515**, 1 (2012).
265. G. Altarelli, R. Barbieri, and S. Jadach, Nucl. Phys. **B369**, 3 (1992) and *ibid.* **B376**, 444(E) (1992).
266. A. De Rújula *et al.*, Nucl. Phys. **B384**, 3 (1992);  
K. Hagiwara *et al.*, Phys. Rev. **D48**, 2182 (1993);  
C.P. Burgess *et al.*, Phys. Rev. **D49**, 6115 (1994);  
Z. Han and W. Skiba, Phys. Rev. **D71**, 075009 (2005);  
G. Cacciapaglia *et al.*, Phys. Rev. **D74**, 033011 (2006);  
A. Falkowski and F. Riva, JHEP **1502**, 039 (2015);  
J. Ellis, V. Sanz and T. You, JHEP **1503**, 157 (2015);  
A. Efrati, A. Falkowski and Y. Soreq, JHEP **1507**, 018 (2015).

## 11. STATUS OF HIGGS BOSON PHYSICS

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## I. Introduction

Understanding the mechanism that breaks the electroweak symmetry and generates the masses of the known elementary particles<sup>1</sup> has been one of the fundamental endeavors in particle physics. The discovery in 2012 by the ATLAS [1] and the CMS [2] Collaborations of a new resonance with a mass of approximately 125 GeV and the subsequent studies of its properties with a much

<sup>1</sup> In the case of neutrinos, it is possible that the electroweak symmetry breaking mechanism plays only a partial role in generating the observed neutrino masses, with additional contributions at a higher scale via the so called see-saw mechanism.

larger data set have provided the first portrait of this mechanism. The mass of this boson has been precisely measured and its production and decay rates are found to be consistent, within errors, with the standard model (SM) predictions. Nevertheless, many theoretical questions remain unanswered and new conundrums about what lies behind the Higgs boson have come to fore. Four years since its discovery, the Higgs boson has turned into a new tool to explore the manifestations of the SM and to probe the physics landscape beyond it.

In the SM [3] the electroweak interactions are described by a gauge field theory invariant under the  $SU(2)_L \times U(1)_Y$  symmetry group. The mechanism of electroweak symmetry breaking (EWSB) [4] provides a general framework to keep untouched the structure of these gauge interactions at high energies and still generate the observed masses of the  $W$  and  $Z$  gauge bosons. The EWSB mechanism posits a self-interacting complex doublet scalar field, whose CP-even neutral component acquires a vacuum expectation value (VEV)  $v \approx 246$  GeV, which sets the scale of electroweak symmetry breaking. Three massless Goldstone bosons are generated and are absorbed to give masses to the  $W$  and  $Z$  gauge bosons. The remaining component of the complex doublet becomes the Higgs boson – a new fundamental scalar particle. The masses of all fermions are also a consequence of EWSB since the Higgs doublet is postulated to couple to the fermions through Yukawa interactions.

The true structure behind the newly discovered boson – including the exact dynamics that triggers the Higgs VEV – and the corresponding ultraviolet completion is, however, still unsolved. Even if the discovered boson has weak couplings to all known SM degrees of freedom, it is not excluded that it is part of an extended symmetry structure or that it emerges from a light resonance of a strongly coupled sector. It needs to be established whether the Higgs boson is solitary or whether other states populate the EWSB sector.

Without the Higgs boson, the calculability of the SM would have been spoiled. In particular, perturbative unitarity [5, 6] would be lost at high energies as the longitudinal  $W/Z$  boson scattering amplitude would grow as the centre-of-mass energy increases. Moreover, the radiative corrections to the gauge boson self-energies would exhibit dangerous logarithmic divergences that would be difficult to reconcile with EW precision data. With the discovery of the Higgs boson, it has been experimentally established that the SM is based on a gauge theory that could a priori be consistently extrapolated to the Planck scale. The Higgs boson must have couplings to  $W/Z$  gauge bosons and fermions precisely as those in the SM to maintain the consistency of the theory at high energies, hence, formally there is no need for new physics at the EW scale. However, as the SM Higgs boson is a scalar particle, and therefore without a symmetry to protect its mass, at the quantum level it has sensitivity to the physics in the ultraviolet. Quite generally, the Higgs mass parameter  $m$  may be affected by the presence of heavy particles. Specifically, in presence of fermion and boson particles with squared masses  $m_i^2 + \lambda_i^2 \phi^2/2$ , the running of the mass parameter from the scale  $\mu$  to the scale  $Q$  reads

$$m^2(Q) = m^2(\mu) + \delta m^2, \quad (11.1)$$

$$\delta m^2 = \sum_i g_i (-1)^{2S_i} \frac{\lambda_i^2 m_i^2}{32\pi^2} \log\left(\frac{Q^2}{\mu^2}\right), \quad (11.2)$$

where the sum is over all particles and  $g_i$  and  $S_i$  correspond to the number of degrees of freedom and the spin of the particle  $i$ . Therefore, particles that couple to the Higgs and have a large squared mass parameter  $m_i^2$  would induce very large corrections to the Higgs mass parameter, demanding a large fine tuning to explain why  $m^2$  remains small. Hence, in general, light scalars like the Higgs boson cannot naturally survive in the presence of heavy states at the grand-unification, string or Planck scales. This is known as the hierarchy or naturalness problem [7].

There are two broad classes of models addressing the naturalness problem<sup>2</sup>: one is based on a new fermion-boson symmetry in nature

called supersymmetry (SUSY) [9–11]. This is a weakly coupled approach to EWSB, and in this case, the Higgs boson remains elementary and the corrections to its mass are screened at the scale at which SUSY is broken and remain insensitive to the details of the physics at higher scales. These theories predict at least three neutral Higgs particles and a pair of charged Higgs particles [12]. One of the neutral Higgs bosons, most often the lightest CP-even Higgs, has properties that resemble those of the SM Higgs boson. It is referred to as a SM-like Higgs boson, meaning that its VEV is predominantly responsible for EWSB, and hence has SM-like couplings to the  $W$  and  $Z$  gauge bosons.

The other approach invokes the existence of strong interactions at a scale of the order of a TeV or above and induces strong breaking of the electroweak symmetry [13]. In the original incarnation of this second approach, dubbed technicolor, the strong interactions themselves trigger EWSB without the need of a Higgs boson. Another possibility, more compatible with the ATLAS and CMS discovery, is that the strong interactions produce four light resonances identified with the Higgs doublet and EWSB proceeds through vacuum misalignment [14] (see Refs. [15, 16] for recent reviews).

Both approaches can have important effects on the phenomenology of the Higgs boson associated with EWSB. Also, in each case the Higgs role in unitarization is shared by other particles: additional Higgs bosons in supersymmetry, or new particles in the strong sector.

A third option has also been considered in the literature. It is also a variation of technicolor or Higgsless models [13, 17]. In light of the Higgs boson discovery these models are ruled out. Nevertheless, there still exists the possibility that the Higgs boson discovered at the LHC is in fact the Goldstone boson of the spontaneous breaking of scale invariance at a scale  $f$  [18, 19]. However, given the good agreement of the coupling measurements with the SM predictions, this dilaton/radion scenario now requires involved model-building.

The naturalness problem has been the prime argument for new physics at the TeV scale, and sizable effects on the Higgs boson properties were expected. But the apparent agreement of the Higgs couplings with the SM predictions, together with the strong bounds inherited from precision electroweak and flavor data leaves open the possibility that the Higgs boson may very well be elementary, weakly coupled and solitary up to the Planck scale. However, absence of evidence is not evidence of absence. It is possible that new states present at the TeV scale to stabilize the Higgs mass might simply be elusive at the LHC because they do not carry a color charge. Twin Higgs [20] models were the first incarnation of this neutral naturalness idea [21]. A more extreme recent proposal [22] relies on the cosmological evolution of the Universe to drive the Higgs boson mass to a value much smaller than the cutoff of the theory and alleviates the hierarchy problem without the need for TeV scale new physics.

Extensions of the SM Higgs sector without low-energy supersymmetry will also be discussed in this review. These type of models do not address the naturalness problem in a specific manner, but provide grounds to explore new Higgs boson signals in a more model-independent way, with different types of coupling structure to fermions and gauge bosons. Extended Higgs sectors are usually quite restricted by experimental constraints from precision electroweak measurements as well as constraints from flavor-changing neutral- and charged-current effects.

Section II is a review of the Higgs boson of the SM, discussing its properties and the production mechanisms and decay rates. In Section III, the SM Higgs boson analysis channels are described. In Section IV, the combination of the main analysis channels is discussed. In Section V, measurements of the main quantum numbers and the total width of the Higgs boson are given. In Section VI, a general theoretical framework to describe the deviations of the Higgs couplings from the SM predictions is introduced and the experimental measurements of these Higgs couplings is reviewed together with the analysis establishing the spin and CP-properties of the Higgs boson. Section VII presents, in detail, some of the most interesting models proposed for Higgs extensions of the SM and considers their experimental signatures. Section VIII provides a brief outlook.

<sup>2</sup> Another solution to the naturalness problem is to lower the fundamental scale of quantum gravity, like for instance in models with large extra-dimensions, see Ref. [8].

## II. The standard model and the mechanism of electroweak symmetry breaking

In the SM [3], electroweak symmetry breaking [4] is responsible for generating mass for the  $W$  and  $Z$  gauge bosons rendering the weak interactions short range. The SM scalar potential reads:

$$V(\Phi) = m^2 \Phi^\dagger \Phi + \lambda (\Phi^\dagger \Phi)^2 \quad (11.3)$$

with the Higgs field  $\Phi$  being a self-interacting  $SU(2)$  complex doublet (four real degrees of freedom) with weak hypercharge  $Y=1$  (the hypercharge is normalized such that  $Q = T_{3L} + Y/2$ ):

$$\Phi = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2}\phi^+ \\ \phi^0 + ia^0 \end{pmatrix}, \quad (11.4)$$

where  $\phi^0$  and  $a^0$  are the CP-even and CP-odd neutral components, and  $\phi^+$  is the complex charged component of the Higgs doublet, respectively.  $V(\Phi)$  is the most general renormalizable scalar potential and if the quadratic term is negative the neutral component of the scalar doublet acquires a non-zero vacuum expectation value (VEV)

$$\langle \Phi \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v \end{pmatrix}, \quad (11.5)$$

with  $\phi^0 = H + \langle \phi^0 \rangle$  and  $\langle \phi^0 \rangle \equiv v$ , inducing the spontaneous breaking of the SM gauge symmetry  $SU(3)_C \times SU(2)_L \times U(1)_Y$  into  $SU(3)_C \times U(1)_{em}$ . The global minimum of the theory defines the ground state, and spontaneous symmetry breaking implies that there is a symmetry of the system (Lagrangian) that is not respected by the ground state. The Higgs field permeates the entire universe and through its self-interactions can cause spontaneous electroweak symmetry breaking (EWSB) in the vacuum. From the four generators of the  $SU(2)_L \times U(1)_Y$  gauge group, three are spontaneously broken, implying that they lead to non-trivial transformations of the ground state and indicate the existence of three massless Goldstone bosons identified with three of the four Higgs field degrees of freedom. The Higgs field couples to the  $W_\mu$  and  $B_\mu$  gauge fields associated with the  $SU(2)_L \times U(1)_Y$  local symmetry through the covariant derivative appearing in the kinetic term of the Higgs Lagrangian,

$$\mathcal{L}_{\text{Higgs}} = (D_\mu \Phi)^\dagger (D^\mu \Phi) - V(\Phi), \quad (11.6)$$

where  $D_\mu \Phi = (\partial_\mu + ig\sigma^a W_\mu^a/2 + ig'Y B_\mu/2)\Phi$ ,  $g$  and  $g'$  are the  $SU(2)$  and  $U(1)$  gauge couplings, respectively, and  $\sigma^a, a = 1, 2, 3$  are the usual Pauli matrices. As a result, the neutral and the two charged massless Goldstone degrees of freedom mix with the gauge fields corresponding to the broken generators of  $SU(2)_L \times U(1)_Y$  and become the longitudinal components of the  $Z$  and  $W$  physical gauge bosons, respectively. The  $Z$  and  $W$  gauge bosons acquire masses,

$$M_W^2 = \frac{g^2 v^2}{4}, \quad M_Z^2 = \frac{(g'^2 + g^2) v^2}{4}. \quad (11.7)$$

The fourth generator remains unbroken since it is the one associated to the conserved  $U(1)_{em}$  gauge symmetry, and its corresponding gauge field, the photon, remains massless. Similarly the eight color gauge bosons, the gluons, corresponding to the conserved  $SU(3)_C$  gauge symmetry with 8 unbroken generators, also remain massless. Hence, from the initial four degrees of freedom of the Higgs field, two are absorbed by the  $W^\pm$  gauge bosons, one by the  $Z$  gauge boson, and there is one remaining degree of freedom,  $H$ , that is the physical Higgs boson — a new scalar particle. The Higgs boson is neutral under the electromagnetic interactions and transforms as a singlet under  $SU(3)_C$  and hence does not couple at tree level to the massless photons and gluons.

The fermions of the SM acquire mass through new renormalizable interactions between the Higgs field and the fermions: the Yukawa interactions,

$$\mathcal{L}_{\text{Yukawa}} = -\hat{h}_{d_{ij}} \bar{q}_{Li} \Phi d_{Rj} - \hat{h}_{u_{ij}} \bar{q}_{Li} \tilde{\Phi} u_{Rj} - \hat{h}_{l_{ij}} \bar{l}_{Li} \Phi e_{Rj} + h.c., \quad (11.8)$$

which respect the symmetries of the SM but generate fermion masses once EWSB occurs. In the above,  $\tilde{\Phi} = i\sigma_2 \Phi^*$  and  $q_L$  ( $l_L$ ) and  $u_R$ ,  $d_R$  ( $e_R$ ) are the quark (lepton)  $SU(2)_L$  doublets and singlets, respectively, while each term is parametrized by a  $3 \times 3$  matrix in family space. The mass term for neutrinos is omitted, but could be added in an analogous manner to the up-type quarks when right-handed neutrinos are supplementing the SM particle content. Once the Higgs acquires a VEV, and after rotation to the fermion mass eigenstate basis that also diagonalizes the Higgs-fermion interactions,  $\hat{h}_{f_{ij}} \rightarrow h_{f_{ij}} \delta_{ij}$ , all fermions acquire a mass given by  $m_{f_i} = h_{f_i} v / \sqrt{2}$ . The indices  $i, j = 1, 2, 3$  refer to the three families in the up-quark, down-quark or charged lepton sectors. It should be noted that the EWSB mechanism provides no additional insight on possible underlying reasons for the large variety of masses of the fermions, often referred to as the flavor hierarchy. The fermion masses, accounting for a large number of the free parameters of the SM, are simply translated into Yukawa couplings  $h_f$ .

### II.1. The SM Higgs boson mass, couplings and quantum numbers

The SM Higgs boson is a CP-even scalar of spin 0. Its mass is given by  $m_H = \sqrt{2\lambda} v$ , where  $\lambda$  is the Higgs self-coupling parameter in  $V(\Phi)$ . The expectation value of the Higgs field,  $v = (\sqrt{2}G_F)^{-1/2} \approx 246$  GeV, is fixed by the Fermi coupling  $G_F$ , which is determined with a precision of 0.6 ppm from muon decay measurements [23]. The quartic coupling  $\lambda$  is a free parameter in the SM, and hence, there is no a priori prediction for the Higgs mass. Moreover the sign of the mass parameter  $m^2 = -\lambda v^2$  is crucial for the EW symmetry breaking to take place, but it is not specified in the SM. The experimentally measured Higgs mass,  $m_H \simeq 125$  GeV, implies that  $\lambda \simeq 0.13$  and  $|m| \simeq 88.8$  GeV. It is interesting to observe that in the SM one needs to assume that the mass term in the potential is negative in order to trigger EWSB. In other theories beyond the SM (BSM), such as supersymmetry, the analogue of the Higgs mass parameter can be made negative dynamically.

The Higgs boson couplings to the fundamental particles are set by their masses. This is a new type of interaction; very weak for light particles, such as up and down quarks, and electrons, but strong for heavy particles such as the  $W$  and  $Z$  bosons and the top quark. More precisely, the SM Higgs couplings to fundamental fermions are linearly proportional to the fermion masses, whereas the couplings to bosons are proportional to the square of the boson masses. The SM Higgs boson couplings to gauge bosons and fermions, as well as the Higgs boson self coupling, are summarized in the following Lagrangian:

$$\mathcal{L} = -g_H f \bar{f} f H + \frac{g_{HHH}}{6} H^3 + \frac{g_{HHHH}}{24} H^4 + \delta_V V_\mu V^\mu \left( g_{HVV} H + \frac{g_{HHVV}}{2} H^2 \right) \quad (11.9)$$

with

$$g_H f \bar{f} = \frac{m_f}{v}, \quad g_{HVV} = \frac{2m_V^2}{v}, \quad g_{HHVV} = \frac{2m_V^2}{v^2}, \quad (11.10)$$

$$g_{HHH} = \frac{3m_H^2}{v}, \quad g_{HHHH} = \frac{3m_H^2}{v^2}, \quad (11.11)$$

where  $V = W^\pm$  or  $Z$  and  $\delta_W = 1, \delta_Z = 1/2$ . As a result, the dominant mechanisms for Higgs boson production and decay involve the coupling of  $H$  to  $W$ ,  $Z$  and/or the third generation quarks and leptons. The Higgs boson coupling to gluons [24, 25] is induced at leading order by a one-loop process in which  $H$  couples to a virtual  $t\bar{t}$  pair. Likewise, the Higgs boson coupling to photons is also generated via loops, although in this case the one-loop graph with a virtual  $W^+W^-$  pair provides the dominant contribution [12] and the one involving a virtual  $t\bar{t}$  pair is subdominant.

### II.2. The SM custodial symmetry

The SM Higgs Lagrangian,  $\mathcal{L}_{\text{Higgs}} + \mathcal{L}_{\text{Yukawa}}$  of Eq. (11.6) and Eq. (11.8), is, by construction,  $SU(2)_L \times U(1)_Y$  gauge invariant, but it also has an approximate global symmetry. In the limit  $g' \rightarrow 0$  and  $h_f \rightarrow 0$ , the Higgs sector has a global  $SU(2)_R$  symmetry, and hence in

such a limit it is invariant under a global  $SU(2)_L \times SU(2)_R$  symmetry, with  $SU(2)_L$  just being the global variant of the SM chiral gauge symmetry. This symmetry is preserved for non-vanishing Yukawa couplings, provided  $h_u = h_d$ . Once the Higgs acquires a VEV, both the  $SU(2)_L$  and  $SU(2)_R$  symmetry groups are broken but the subgroup  $SU(2)_{L+R}$  remains unbroken and is the subgroup that defines the custodial symmetry of the SM [26].

In the limit  $g' \rightarrow 0$ , the  $W$  and  $Z$  gauge bosons have equal mass and form a triplet of the  $SU(2)_{L+R}$  unbroken global symmetry. Using the expressions for the  $W$  and  $Z$  gauge boson masses in term of the gauge couplings, one obtains

$$\frac{M_W^2}{M_Z^2} = \frac{g^2}{g'^2 + g^2} = \cos^2 \theta_W \quad \text{or} \quad \rho \equiv \frac{M_W^2}{M_Z^2 \cos^2 \theta_W} = 1 \quad (11.12)$$

at tree level. The custodial symmetry protects the above relation between the  $W$  and  $Z$  masses under radiative corrections. All corrections to the  $\rho$  parameter are therefore proportional to terms that break the custodial symmetry. For instance, radiative corrections involving the Higgs are proportional to  $g'^2$ . Since  $m_t \neq m_b$ , there are also relevant radiative corrections generated by massive fermions. They are proportional to  $m_t^2 + m_b^2 - 2(m_t^2 m_b^2) \log(m_t^2/m_b^2)/(m_t^2 - m_b^2)$  [27].

One can conceive of BSM theories in which the Higgs is a pseudo Nambu–Goldstone boson of a strongly interacting sector [28], and/or where there are additional degrees of freedom that may contribute to the  $W$  and  $Z$  mass via virtual loops, but in as much as the electroweak sector has a manifest custodial symmetry, the theory is protected from large radiative corrections. Precision measurements of electroweak observables are powerful in constraining such large radiative corrections. The custodial isospin symmetry is also a powerful probe of BSM physics. For a pedagogical discussion, see Ref. [29].

energy evolution of  $\lambda$  shows that it becomes negative at energies  $\Lambda = \mathcal{O}(10^{10} - 10^{12})$  GeV, with a broader range if the top quark mass exceeds its current measured value by  $3\sigma$ . When this occurs, the SM Higgs potential develops an instability and the long term existence of the EW vacuum is challenged. This behavior may call for new physics at an intermediate scale before the instability develops, i.e., below  $M_{Planck}$  or, otherwise, the electroweak vacuum remains metastable [32]. Reference [33] studied how new physics at  $M_{Planck}$  could influence the stability of the EW vacuum and possibly modify this conclusion. The consequences of the instability of the EW vacuum on high-scale inflation have been discussed in Refs. [34].

Within the SM framework, the relevant question is related to the lifetime of the EW metastable vacuum that is determined by the rate of quantum tunneling from this vacuum into the true vacuum of the theory. The running of the Higgs self coupling slows down at high energies with a cancellation of its  $\beta$ -function at energies just one to two orders of magnitude below the Planck scale [31, 35]. This slow evolution of the quartic coupling is responsible for saving the EW vacuum from premature collapse, allowing it to survive much longer times than those relevant from astrophysical considerations. It might help the Higgs boson to play the role of an inflaton [36] (see, however, Ref. [37] and references therein for potential issues with this Higgs-as-inflaton idea).

## II.4. Higgs production and decay mechanisms

Reviews of the SM Higgs boson's properties and phenomenology, with an emphasis on the impact of loop corrections to the Higgs boson decay rates and cross sections, can be found in Refs. [38–45]. The state-of-the-art of the theoretical calculations in the main different production channels is summarized in Table 11.1.

**Table 11.1:** State-of-the-art of the theoretical calculations in the main different Higgs production channels in the SM, and main MC tools used in the simulations

ggF	VBF	VH	ttH
Fixed order:	Fixed order:	Fixed order:	Fixed order:
NNLO QCD + NLO EW	NNLO QCD	NLO QCD+EW	NLO QCD
(HIGLU, iHixs, FeHiPro, HNNLO)	(VBF@NNLO)	(V2HV and HAWK)	(Powheg)
Resummed:	Fixed order:	Fixed order:	(MG5_aMC@NLO)
NNLO + NNLL QCD	NLO QCD + NLO EW	NNLO QCD	
(HRes)	(HAWK)	(VH@NNLO)	
Higgs $p_T$ :			
NNLO+NNLL			
(HqT, HRes)			
Jet Veto:			
N3LO+NNLL			

### II.3. Stability of the Higgs potential

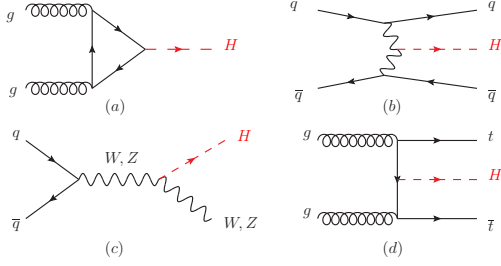
The discovery of a scalar particle with mass  $m_H \approx 125$  GeV has far reaching consequences within the SM framework. In particular, the precise value of  $m_H$  determines the value of the quartic coupling  $\lambda$  at the electroweak scale and makes it possible to study its behavior up to high energy scales. A larger value of  $m_H$  would have implied that the Higgs self-coupling would become non-perturbative at some scale  $\Lambda$  that could be well below the Planck scale. Specifically, from the measured values of the Higgs mass, the top-quark mass, the  $W$  and  $Z$  boson masses, and the strong gauge coupling, all within their experimental uncertainties, it follows that, as with the SM gauge and Yukawa couplings, the Higgs quartic coupling remains perturbative all the way up to  $M_{Planck}$  [5, 6, 30], thereby rendering the SM a consistent, calculable theory.

The recently measured Higgs mass, however, generates an EW Higgs potential in which the vacuum state is at the edge between being stable and metastable. Indeed, allowing all relevant SM observables to fluctuate within their experimental and theoretical uncertainties, the metastability condition seems to be favored [31]. The high

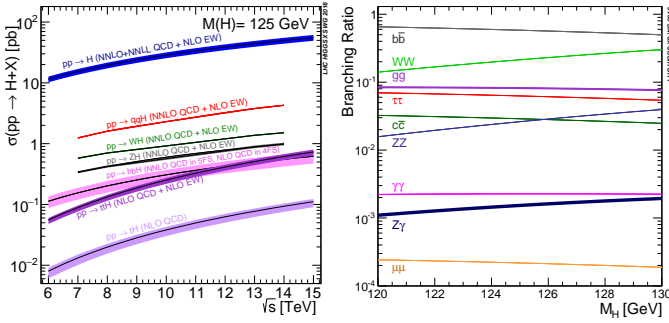
#### II.4.1. Production mechanisms at hadron colliders

The main production mechanisms at the Tevatron collider and the LHC are gluon fusion, weak-boson fusion, associated production with a gauge boson and associated production with a pair of top/antitop quarks. Figure 11.1 depicts representative diagrams for these dominant Higgs production processes.

The cross sections for the production of a SM Higgs boson as a function of  $\sqrt{s}$ , the center of mass energy, for  $pp$  collisions, including bands indicating the theoretical uncertainties, are summarized in Fig. 11.2(left) [46]. A detailed discussion, including uncertainties in the theoretical calculations due to missing higher-order effects and experimental uncertainties on the determination of SM parameters involved in the calculations can be found in Refs. [42–45]. These references also contain state-of-the-art discussions on the impact of PDF's uncertainties, QCD scale uncertainties and uncertainties due to different matching procedures when including higher-order corrections matched to parton shower simulations as well as uncertainties due to hadronization and parton-shower events.



**Figure 11.1:** Generic Feynman diagrams contributing to the Higgs production in (a) gluon fusion, (b) weak-boson fusion, (c) Higgs-strahlung (or associated production with a gauge boson) and (d) associated production with top quarks.



**Figure 11.2:** (Left) The SM Higgs boson production cross sections as a function of the center of mass energy,  $\sqrt{s}$ , for  $pp$  collisions. The theoretical uncertainties [46] are indicated as bands. (Right) The branching ratios for the main decays of the SM Higgs boson near  $m_H = 125$  GeV. The theoretical uncertainties [44, 45] are indicated as bands.

**Table 11.2:** The SM Higgs boson production cross sections for  $m_H = 125$  GeV in  $pp$  collisions ( $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for the Tevatron), as a function of the center of mass energy,  $\sqrt{s}$ . The predictions for the LHC energies are taken from Refs. [42–45], the ones for the Tevatron energy are from Ref. [47]. The predictions for the  $ggF$  channel do not include the latest N3LO results which significantly reduce the theoretical uncertainties.

$\sqrt{s}$ (TeV)	Production cross section (in pb) for $m_H = 125$ GeV					total
	ggF	VBF	WH	ZH	$t\bar{t}H$	
1.96	$0.95^{+17\%}_{-17\%}$	$0.065^{+8\%}_{-7\%}$	$0.13^{+8\%}_{-8\%}$	$0.079^{+8\%}_{-8\%}$	$0.004^{+10\%}_{-10\%}$	1.23
7	$15.3^{+10\%}_{-10\%}$	$1.24^{+2\%}_{-2\%}$	$0.58^{+3\%}_{-3\%}$	$0.34^{+4\%}_{-4\%}$	$0.09^{+8\%}_{-14\%}$	17.5
8	$19.5^{+10\%}_{-11\%}$	$1.60^{+2\%}_{-2\%}$	$0.70^{+3\%}_{-3\%}$	$0.42^{+5\%}_{-5\%}$	$0.13^{+8\%}_{-13\%}$	22.3
13	$44.1^{+11\%}_{-11\%}$	$3.78^{+2\%}_{-2\%}$	$1.37^{+2\%}_{-2\%}$	$0.88^{+5\%}_{-5\%}$	$0.51^{+9\%}_{-13\%}$	50.6
14	$49.7^{+11\%}_{-11\%}$	$4.28^{+2\%}_{-2\%}$	$1.51^{+2\%}_{-2\%}$	$0.99^{+5\%}_{-5\%}$	$0.61^{+9\%}_{-13\%}$	57.1

Table 11.2, from Refs. [42–45], summarizes the Higgs boson production cross sections and relative uncertainties for a Higgs mass of 125 GeV, for  $\sqrt{s} = 7, 8, 13$  and 14 TeV. The Higgs boson production cross sections in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for the Tevatron are obtained from Ref. [47].

### (i) Gluon fusion production mechanism

At high-energy hadron colliders, the Higgs boson production mechanism with the largest cross section is the gluon-fusion process,  $gg \rightarrow H + X$ , mediated by the exchange of a virtual, heavy top quark [48]. Contributions from lighter quarks propagating in the loop are suppressed proportional to  $m_q^2$ . QCD radiative corrections to the gluon-fusion process are very important and have been studied in detail. Including the full dependence on the (top, bottom, charm) quark and Higgs boson masses, the cross section has been calculated

at the next-to-leading order (NLO) in  $\alpha_s$  [49, 50]. To a very good approximation, the leading top-quark contribution can be evaluated in the limit  $m_t \rightarrow \infty$  by matching the SM to an effective theory. The gluon-fusion amplitude is then evaluated from an effective Lagrangian containing a local  $HG_{\mu\nu}^a G^{a\mu\nu}$  operator [24, 25]. In this approximation the cross section is known at NLO [51], at next-to-next-to-leading order (NNLO) [52], and recently the computation at next-to-next-to-next-to-leading order (N3LO) has been completed [53]. The validity of the large top-quark mass approximation in NNLO calculations has been established at the percent level by means of approximate calculations of the  $m_t$  dependence based on asymptotic expansions [54]. Moreover, the validity of the effective theory with infinite  $m_t$  is greatly enhanced by rescaling the result by the exact LO result:  $\sigma = (\sigma_{m_t}^{LO} / \sigma_{m_t=\infty}^{LO}) \times \sigma_{m_t=\infty}$  [45].

The NLO QCD corrections increase the leading-order prediction for the cross section by about 80%, the NNLO corrections further enhance the cross section by approximately 30% (at  $\mu_f = \mu_r = m_H/2$ ). Electroweak radiative corrections have been computed at NLO and increase the cross section by about 5% for  $m_H \simeq 125$  GeV [55]. Mixed QCD-electroweak corrections of  $O(\alpha_s)$  have been calculated in Ref. [56].

The NLO and NNLO fixed-order QCD predictions for the gluon-fusion cross section have been improved by resumming the soft, virtual and collinear gluon contributions to the cross section at next-to-next-to-leading logarithmic (NNLL) and partial NNNLL accuracy [57]. Precise predictions for the gluon-fusion cross section for different Higgs boson masses and LHC energies, and including detailed error estimates, have been obtained by combining the NNLO fixed-order QCD results with soft-gluon resummation at NNLL or NNNLL accuracy and two-loop electroweak corrections, and using the most recent sets of parton distribution functions [56, 58].

The perturbative QCD computation has been recently extended to N3LO. At this order the perturbation series is rather stable with a mere enhancement of 3% only, with a central value completely insensitive to threshold resummation effects with the scale choice mentioned above [53, 59, 45]. At the LHC with a center-of-mass energy of 13 TeV, the most up-to-date value for the production cross section of a 125 GeV Higgs boson amounts to [45]

$$\sigma_{ggF}^{N3LO} = 48.6 \text{ pb}^{+2.2 \text{ pb}(+4.6\%)}_{-3.3 \text{ pb}(-6.7\%)}(\text{theory}) \pm 1.6 \text{ pb}(3.2\%)(\text{PDF} + \alpha_s).$$

The difference between this result and the value quoted in Table 11.2 is due to several effects that include: the choice of optimal renormalization and factorization scales, the effect of the N3LO corrections, the different sets of parton distribution functions and value of  $\alpha_s$ , as well as smaller differences due to the treatment of finite quark-mass effects [53].

Besides considering the inclusive Higgs boson production cross section at the LHC, it is important to study differential distributions in order to probe the properties of the Higgs boson in a detailed way. A more exclusive account of Higgs production is also required because experimental analyses often impose cuts on the final states in order to improve the signal-to-background ratio. To this end, it is useful to define benchmark cuts and compare the differential distributions obtained at various levels of theoretical accuracy (i.e., at NLO or NNLO) with Monte Carlo generators. Many search modes for the Higgs boson are carried out by separating the events according to the number of jets or the transverse momentum and rapidity of the Higgs boson. For  $p_T < 30$  GeV, predictions for the transverse-momentum distribution can only be trusted after large logarithms of the form  $\alpha_s^n \ln^{2n-1}(m_H/p_T)$  have been resummed to all orders in perturbation theory [60]. This has been accomplished with NNLL accuracy [61], and the results have been matched onto the fixed-order prediction at NNLO [62]. Electroweak corrections have been studied in Ref. [63]. The effect of the non-zero quark mass on the  $p_T$  spectrum has been considered in Refs. [64, 65], while the effect of the finite top mass on other differential observables has been studied in Refs. [66, 67]. There has been much activity in computing Higgs plus jet(s) production processes at NLO (see e.g. Refs. [68] and [69] for associated production with one and two jets, respectively), and even at NNLO [70]. In addition, efforts to improve the calculation of the Higgs production



cross section with a jet veto (the “0-jet bin”) by resumming large logarithms of the form  $\alpha_s^n \ln^{2n-1}(m_H/p_T^{\text{veto}})$  at NNLL order and beyond [71] have been made. Recently, reference results for the resummed cross section at NNLL have been combined with the N3LO result for the inclusive cross section to obtain accurate predictions for the jet-veto efficiency and zero-jet cross section [72]. Accurate predictions for the jet-veto cross section are required, e.g., to suppress the  $t\bar{t}$  background in the  $H \rightarrow WW$  channel [73].

### (ii) Vector boson fusion production mechanism

The SM Higgs production mode with the second-largest cross section at the LHC is vector boson fusion (VBF). At the Tevatron collider, VBF also occurred, but for  $m_H = 125$  GeV had a smaller cross section than Higgs production in association with a  $W$  or  $Z$  boson. Higgs production via VBF,  $qq \rightarrow qqH$ , proceeds by the scattering of two (anti-)quarks, mediated by  $t$ - or  $u$ -channel exchange of a  $W$  or  $Z$  boson, with the Higgs boson radiated off the weak-boson propagator. The scattered quarks give rise to two hard jets in the forward and backward regions of the detector. Because of the color-singlet nature of the weak-gauge boson exchange, gluon radiation from the central-rapidity regions is strongly suppressed. These characteristic features of VBF processes can be exploited to distinguish them from overwhelming QCD backgrounds, including gluon-fusion induced Higgs + 2 jet production, and from  $s$ -channel  $WH$  or  $ZH$  production with a hadronically decaying weak gauge boson. After the application of specific selection cuts, the VBF channel provides a particularly clean environment, not only for Higgs searches but also for the determination of Higgs boson couplings at the LHC [74].

Computations for total cross sections and differential distributions to Higgs production via VBF including NLO QCD and EW corrections have been presented in Refs. [39, 75] and are available in the form of flexible parton-level Monte-Carlo generators. Parton-shower effects have been considered in Ref. [76]. The NNLO QCD corrections to the total rate have been presented in Refs. [77]. They reduce the residual scale uncertainties on the inclusive cross section to approximately 2%. The uncertainties due to parton distributions are estimated to be at the same level. Fully differential predictions at NNLO have been computed recently [78], suggesting that the cross section under VBF cuts receives NNLO corrections that are larger than in the inclusive case and may reach O(5–6%).

### (iii) $WH$ and $ZH$ associated production mechanism

The next most relevant Higgs boson production mechanisms after gluon fusion and VBF at the LHC, and the most relevant ones after gluon fusion at the Tevatron collider, are associated production with  $W$  and  $Z$  gauge bosons. The cross sections for the associated production processes,  $pp \rightarrow VH + X$ , with  $V = W^\pm, Z$  receive contributions at NLO given by NLO QCD corrections to the Drell–Yan cross section [79–81] and from NLO EW corrections. The latter, unlike the QCD corrections, do not respect the factorization into Drell–Yan production since there are irreducible box contributions already at one loop [82]. At NNLO, the Drell–Yan-like corrections to  $WH$  production also give the bulk of the corrections to  $ZH$  production [83]. For  $ZH$  production there are, however, gluon-gluon induced contributions that do not involve a virtual  $Z$  gauge boson but are such that the  $Z$  gauge boson and  $H$  boson couple to gluons via top-quark loops [84]. In addition,  $WH$  and  $ZH$  production receive non Drell–Yan-like corrections in the  $q\bar{q}'$  and  $q\bar{q}$  initiated channels, respectively, at the NNLO level, where the Higgs is radiated off top-quark loops [85]. The full QCD corrections up to NNLO order, the NLO EW corrections and the NLO corrections to the gluon-gluon channel are available in VH@NNLO [86].

As neither the Higgs boson nor the weak gauge bosons are stable particles, their decays also have to be taken into account. Providing full kinematical information for the decay products can furthermore help in the suppression of large QCD backgrounds. Differential distributions for the processes  $pp \rightarrow WH \rightarrow \nu_\ell \bar{\nu}_\ell H$  and  $pp \rightarrow ZH \rightarrow \ell^+ \ell^- H / \nu_\ell \bar{\nu}_\ell H$ , including NLO QCD and EW corrections, have been presented in Ref. [87]. The NNLO QCD corrections to differential observables for  $WH$  production at the LHC, including the leptonic decays of the  $W$  boson and the decay of the Higgs boson

into a  $b\bar{b}$  pair, are presented in Ref. [88]. Calculations at the same level, including also the  $ZH$  process have been performed [89, 90]. The  $WH$  production mode has also been matched to a parton shower at NNLO accuracy [91]. The  $WH$  and  $ZH$  production modes, together with Higgs production in association with a top-quark pair, provide a relatively clean environment for studying the decay of the Higgs boson into bottom quarks.

### (iv) Higgs production in association with $t\bar{t}$

Higgs radiation off top quarks,  $pp \rightarrow t\bar{t}H$ , can provide important information on the top-Higgs Yukawa coupling and gives access to the Higgs decay into bottom quarks. The LO cross section for this production process was computed in Ref. [92]. Later, the NLO QCD corrections [93] were evaluated yielding a moderate increase in the total cross section of at most 20%, but reducing significantly the scale dependence of the inclusive cross section. The total theoretical errors, estimated by combining the uncertainties from factorization and renormalization scales, strong gauge coupling, and parton distributions, amount to 10–15% of the corresponding inclusive cross section. Interfaces between NLO QCD calculations for  $t\bar{t}H$  production with parton-shower Monte Carlo programs have been provided in Ref. [94]. These programs provide the most flexible tools to date for the computation of differential distributions, including experimental selection cuts and vetoes on the final-state particles and their decay products.

### (v) Other single Higgs production mechanisms at the LHC

The Higgs production in association with a single top quark, though subdominant, can bring valuable information, in particular regarding the sign of the top Yukawa coupling. This is due to an almost totally destructive interference between two large contributions, one where the Higgs couples to a space-like  $W$  boson and the other where it couples to the top quark. This process has been computed at NLO in a five-flavor scheme [95] and amounts to about 90 fb at  $\sqrt{s} = 14$  TeV (with the opposite sign of the top Yukawa coupling, the cross section increases by one order of magnitude).

The Higgs boson production in association with bottom quarks is known at NNLO in the case of five quark flavors [96–98]. The coupling of the Higgs boson to a  $b$  quark is suppressed in the SM by the bottom-quark mass over the Higgs VEV,  $m_b/v$ , implying that associated production of a SM Higgs boson with  $b$  quarks is small at the LHC. Yet, at high energy, large logarithms are present and need to be resummed, leading to an enhancement of the inclusive cross section. At  $\sqrt{s} = 14$  TeV the  $b\bar{b}H$  cross section can be as large as 600 fb, still two orders of magnitude below the ggF production cross section. In a two Higgs doublet model or a supersymmetric model, which will be discussed in Section VII, this coupling is proportional to the ratio of neutral Higgs boson vacuum expectation values,  $\tan\beta$ , and can be significantly enhanced for large values of this ratio. Consequently, the  $b\bar{b}H$  mode can even become the dominant production process for the Higgs boson.

The Higgs production in association with charm quarks is also known at NNLO and is of the order of 85 fb at  $\sqrt{s} = 13$  TeV.

### (vi) Double Higgs production at the LHC

The main interest in the double Higgs production is that it provides invaluable information on the Higgs potential. In particular, it gives access to the Higgs cubic self-interaction. The dominant production is via gluon fusion  $gg \rightarrow HH$ . The NLO [99] and NNLO [100] fixed order corrections to  $gg \rightarrow HH$  are known in the infinite top mass limit and, recently, the complete NLO corrections with all top quark mass effects also became available [101]. The QCD corrections are large, typically doubling the cross section from LO to NLO and further enhancing it by 20% from NLO to NNLO. At the differential level, the destructive interference between the box and the triangle contributions complicates the predictions made in the infinite top mass limit for both the  $HH$  invariant mass and the leading Higgs  $p_T$  distributions. With an inclusive cross section of about 40 fb at  $\sqrt{s} = 13$  TeV and a difficult signal vs. background discrimination, the double Higgs production remains a challenging channel to probe and will greatly benefit from the high-luminosity run of the LHC.

### II.4.2. Production mechanisms at $e^+e^-$ colliders

The main Higgs boson production cross sections at an  $e^+e^-$  collider are the Higgs-strahlung process  $e^+e^- \rightarrow ZH$  [6, 24, 102], and the  $WW$  fusion process [103]  $e^+e^- \rightarrow \bar{\nu}_e \nu_e W^* W^* \rightarrow \bar{\nu}_e \nu_e H$ . The cross-section for the Higgs-strahlung process scales as  $s^{-1}$  and is dominant at low energies, while the cross-section for the  $WW$  fusion process scales as  $\ln(s/m_H^2)$  and dominates at high energies [104–106]. The  $ZZ$  fusion mechanism,  $e^+e^- \rightarrow e^+e^- Z^* Z^* \rightarrow e^+e^- H$ , also contributes to Higgs boson production, with a cross-section suppressed by an order of magnitude with respect to that of  $WW$  fusion. The process  $e^+e^- \rightarrow t\bar{t}H$  [107, 108] becomes important for  $\sqrt{s} \geq 500$  GeV. For a more detailed discussion of Higgs production properties at lepton colliders see, for example, Refs. [40, 41, 109, 110] and references therein.

### II.4.3. SM Higgs branching ratios and total width

For the understanding and interpretation of the experimental results, the computation of all relevant Higgs decay widths is essential, including an estimate of their uncertainties and, when appropriate, the effects of Higgs decays into off-shell particles with successive decays into lighter SM ones. A Higgs mass of about 125 GeV provides an excellent opportunity to explore the Higgs couplings to many SM particles. In particular the dominant decay modes are  $H \rightarrow b\bar{b}$  and  $H \rightarrow WW^*$ , followed by  $H \rightarrow gg$ ,  $H \rightarrow \tau^+\tau^-$ ,  $H \rightarrow c\bar{c}$  and  $H \rightarrow ZZ^*$ . With much smaller rates follow the Higgs decays into  $H \rightarrow \gamma\gamma$ ,  $H \rightarrow \gamma Z$  and  $H \rightarrow \mu^+\mu^-$ . Since the decays into gluons, diphotons and  $Z\gamma$  are loop induced, they provide indirect information on the Higgs couplings to  $WW$ ,  $ZZ$  and  $t\bar{t}$  in different combinations. The uncertainties in the branching ratios include the missing higher-order corrections in the theoretical calculations as well as the errors in the SM input parameters, in particular fermion masses and the QCD gauge coupling, involved in the decay. In the following the state-of-the-art of the theoretical calculations will be discussed and the reader is referred to Refs. [42, 43, 111] for detail.

The evaluation of the radiative corrections to the fermionic decays of the SM Higgs are implemented in HDECAY [112] at different levels of accuracy. The computations of the  $H \rightarrow b\bar{b}$  and  $H \rightarrow c\bar{c}$  decays include the complete massless QCD corrections up to N4LO, with a corresponding scale dependence of about 0.1% [113]. Both the electroweak corrections to  $H \rightarrow b\bar{b}$ ,  $c\bar{c}$  as well as  $H \rightarrow \tau^+\tau^-$  are known at NLO [114] providing predictions with an overall accuracy of about 1–2% for  $m_H \simeq 125$  GeV.

The loop induced decays of the SM Higgs are known at NLO and partially beyond that approximation. For  $H \rightarrow gg$ , the QCD corrections are known up to N3LO in the limit of heavy top quarks [115, 50] and the uncertainty from the scale dependence is about 3%. For the  $H \rightarrow \gamma\gamma$ , the full NLO QCD corrections are available [50, 116] and the three-loop QCD corrections have also been evaluated [117]. The NLO electroweak corrections to  $H \rightarrow gg$  and  $H \rightarrow \gamma\gamma$  have been computed in Ref. [118]. All these corrections are implemented in HDECAY [112]. For  $m_H = 125$  GeV, the overall impact of known QCD and EW radiative effects turns out to be well below 1%. In addition, the contribution of the  $H \rightarrow \gamma e^+e^-$  decay via virtual photon conversion has been computed in Ref. [119]. The partial decay width  $H \rightarrow Z\gamma$  is only implemented at LO in HDECAY, including the virtual  $W$ , top-, bottom-, and  $\tau$ -loop contributions. The QCD corrections have been calculated and are at the percent level [120]. The theoretical uncertainty due to unknown electroweak corrections is estimated to be less than 5%, an accuracy that will be hard to achieve in measurements of this processes at the LHC.

The decays  $H \rightarrow WW/ZZ \rightarrow 4f$  can be simulated with the Prophecy4f Monte-Carlo generator [121] that includes complete NLO QCD and EW corrections for Higgs decays into any possible four-fermion final state. All calculations are consistently performed with off-shell gauge bosons, without any on-shell approximation. For the SM Higgs boson the missing higher-order corrections are estimated to be roughly 0.5%. Such uncertainties will have to be combined with the parametric uncertainties, in particular those associated to the bottom-quark mass and the strong gauge coupling, to arrive at the full theory uncertainties. A detailed treatment of the differential distributions for a Higgs decay into four charged leptons in the final state is discussed in Refs. [44, 122].

The total width of a 125 GeV SM Higgs boson is  $\Gamma_H = 4.07 \times 10^{-3}$  GeV, with a relative uncertainty of  $^{+4.0\%}_{-3.9\%}$ . The branching ratios for the most relevant decay modes of the SM Higgs boson as a function of  $m_H$ , including the most recent theoretical uncertainties, are shown in Fig. 11.2(right) and listed for  $m_H = 125$  GeV in Table 11.3. Further details of these calculations can be found in Refs. [111, 123] and in the reviews [39–45].

**Table 11.3:** The branching ratios and the relative uncertainty [44, 45] for a SM Higgs boson with  $m_H = 125$  GeV.

Decay channel	Branching ratio	Rel. uncertainty
$H \rightarrow \gamma\gamma$	$2.27 \times 10^{-3}$	$^{+5.0\%}_{-4.9\%}$
$H \rightarrow ZZ$	$2.62 \times 10^{-2}$	$^{+4.3\%}_{-4.1\%}$
$H \rightarrow W^+W^-$	$2.14 \times 10^{-1}$	$^{+4.3\%}_{-4.2\%}$
$H \rightarrow \tau^+\tau^-$	$6.27 \times 10^{-2}$	$^{+5.7\%}_{-5.7\%}$
$H \rightarrow b\bar{b}$	$5.84 \times 10^{-1}$	$^{+3.2\%}_{-3.3\%}$
$H \rightarrow Z\gamma$	$1.53 \times 10^{-3}$	$^{+9.0\%}_{-8.9\%}$
$H \rightarrow \mu^+\mu^-$	$2.18 \times 10^{-4}$	$^{+6.0\%}_{-5.9\%}$

## III. The experimental profile of the Higgs boson

An indirect experimental bound on the SM Higgs boson can be obtained by comparing precision electroweak data with SM predictions, that have a weak, logarithmic dependence on  $M_H$ . A global fit to electroweak data suggests  $m_H = 96^{+22}_{-19}$  GeV, or  $m_H < 134$  GeV at 90% confidence level [124].

The announcement on July 4, 2012 of the observation [1, 2] at the LHC of a narrow resonance with a mass of about 125 GeV was an important landmark in the decades-long direct search [125, 126] for the SM Higgs boson. Even as this discovery was being announced, ATLAS and CMS continued to accumulate pp collision data at  $\sqrt{s} = 8$  TeV recording a total of about 20 fb $^{-1}$  each at this energy. This data set together with about 5 fb $^{-1}$  recorded at  $\sqrt{s} = 7$  TeV comprised the LHC Run 1  $pp$  collision data set. In the remainder of this section the focus will be on the final results on the measurements of Higgs boson properties with the LHC Run 1 data.

### III.1. The principal discovery channels

For a given  $m_H$ , the sensitivity of a search channel depends on the production cross section of the Higgs boson, its decay branching fraction, reconstructed mass resolution, selection efficiency and the level of background in the final state. For a low-mass Higgs boson ( $110 < m_H(\text{GeV}) < 150$ ) where the natural width is only a few MeV, the five decay channels that play an important role at the LHC are listed in Table 11.4. In the  $H \rightarrow \gamma\gamma$  and  $H \rightarrow ZZ \rightarrow 4\ell$  channels, all final state particles can be very precisely measured and the reconstructed  $m_H$  resolution is excellent. While the  $H \rightarrow W^+W^- \rightarrow \ell^+\nu_\ell \ell^-\bar{\nu}_\ell$  channel has relatively large branching fraction, the  $m_H$  resolution is poor due to the presence of neutrinos. The  $H \rightarrow b\bar{b}$  and the  $H \rightarrow \tau^+\tau^-$  channels suffer from large backgrounds and a poor mass resolution. For  $m_H > 150$  GeV, the sensitive search channels are  $H \rightarrow WW$  and  $H \rightarrow ZZ$  where the  $W$  or  $Z$  boson decays into a variety of leptonic and hadronic final states. These decay channels of the Higgs boson are searched for in the five Higgs boson production processes (ggF, VBF, WH, ZH and  $t\bar{t}H$ ) described in Section II.4.1.

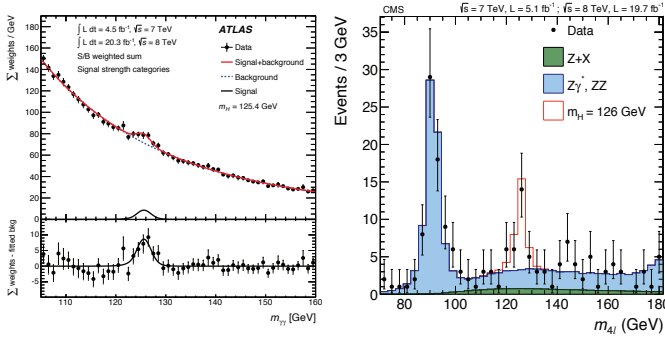
The candidate events in each Higgs boson decay channel are split into several mutually exclusive categories (or event tags) based on the specific topological, kinematic or other features present in the candidate event. The categorization of events increases the sensitivity of the overall analysis and allows a separation of different Higgs

**Table 11.4:** The five principal decay channels for low mass SM Higgs boson searches at the LHC. The numbers reported are for  $m_H = 125$  GeV.

Decay channel	Mass resolution
$H \rightarrow \gamma\gamma$	1–2%
$H \rightarrow ZZ \rightarrow \ell^+\ell^-\ell'^+\ell'^-$	1–2%
$H \rightarrow W^+W^- \rightarrow \ell^+\nu_\ell\ell'^-\bar{\nu}_{\ell'}$	20%
$H \rightarrow b\bar{b}$	10%
$H \rightarrow \tau^+\tau^-$	15%

boson production processes. Most categories are dominated by signal from one Higgs decay mode but contain an admixture of various Higgs production processes. For example, a typical VBF category contains Higgs boson candidates accompanied by two energetic jets ( $\geq 30$  GeV) with a large dijet mass ( $\geq 400$  GeV) and separated by a large pseudorapidity ( $\Delta\eta_{jj} \geq 3.5$ ). While such a category is enriched in Higgs bosons produced via VBF, the contamination from the gluon fusion production mechanism can be significant. Hence a measurement of the signal rate in the VBF category does not imply a measurement of VBF production cross-section. Simulations are used to determine the relative contributions of the various Higgs production modes in a particular category.

### III.1.1. $H \rightarrow \gamma\gamma$



**Figure 11.3:** (Left) The invariant mass distribution of diphoton candidates, with each event weighted by the ratio of signal-to-background in each event category, observed by ATLAS [127]. The residuals of the data with respect to the fitted background are displayed in the lower panel. (Right) The combined  $m_{4\ell}$  distribution from CMS [128] Run 1 data.

In the  $H \rightarrow \gamma\gamma$  channel a search is performed for a narrow peak over a smoothly falling background in the invariant mass distribution of two high  $p_T$  photons. The background in this channel is conspicuous and stems from prompt  $\gamma\gamma$ ,  $\gamma$ +jet and dijet processes. In order to optimize search sensitivity and also to separate the various Higgs production modes, ATLAS and CMS experiments split events into several mutually exclusive categories. Diphoton events containing a high  $p_T$  muon or electron, or missing energy ( $E_T^{\text{miss}}$ ) consistent with the decay of a  $W$  or  $Z$  boson are tagged in the VH production category. Diphoton events containing energetic dijets with a large mass and pseudorapidity difference are assigned to the VBF production category, and the remaining events are considered either in the VH category when the two jets are compatible with the hadronic decay of a  $W$  or a  $Z$ , or in the gluon fusion production category. While the leptonic VH category is relatively pure, the VBF category has significant contamination from the gluon fusion process. A summary of all categories used in this channel is given in Section IV.1 and in Table 11.9. Events which are not picked by any of the above selections are further categorized according to their expected  $m_{\gamma\gamma}$  resolution and signal-to-background ratio. Categories with good  $m_H$  resolution and larger signal-to-background ratio contribute most to the sensitivity of the search.

Both ATLAS and CMS have studied in detail the calibration of the energy response of photons, in particular using  $Z \rightarrow e^+e^-$ ,  $Z \rightarrow \mu^+\mu^-$  and the response of muons in the calorimeter (for ATLAS) from  $Z \rightarrow \mu^+\mu^-$  events. This information is used to correct the fully simulated signal mass lineshapes. In each category, parametric signal models are adjusted to these lineshape to provide a functional form for the signal. Simple monotonic functional forms of the backgrounds are determined by a fit to the full  $m_{\gamma\gamma}$  distribution in each category. All categories are fitted simultaneously to determine the signal yield at a particular mass. In the full dataset, the  $m_{\gamma\gamma}$  distribution after combining all categories is shown for the ATLAS experiment in Fig. 11.3. ATLAS observes [127] an excess over background at  $m_H = 125.4$  GeV with a local significance of  $5.2\sigma$  compared with  $4.6\sigma$  expected for SM Higgs boson at that mass. CMS observes [129] its largest excess at  $m_H = 124.7$  GeV with a local significance of  $5.7\sigma$  compared with  $5.2\sigma$  expected for a SM Higgs boson of that mass.

The signal strength  $\mu = (\sigma \cdot \text{BR})_{\text{obs}}/(\sigma \cdot \text{BR})_{\text{SM}}$ , which is the observed product of the Higgs boson production cross section ( $\sigma$ ) and its branching ratio (BR) in units of the corresponding SM values, is  $1.17 \pm 0.27$  for ATLAS and  $0.78^{+0.26}_{-0.23}$  for CMS at  $m_H = 125.4$  and  $124.7$  GeV, respectively.

### III.1.2. $H \rightarrow ZZ^* \rightarrow \ell^+\ell^-\ell'^+\ell'^-$

In the  $H \rightarrow ZZ^* \rightarrow \ell^+\ell^-\ell'^+\ell'^-$  channel a search is performed for a narrow mass peak over a small continuous background dominated by non-resonant  $ZZ^*$  production from  $q\bar{q}$  annihilation and  $gg$  fusion processes. The contribution and the shape of this background is taken from simulation. The subdominant and reducible backgrounds stem from  $Z + b\bar{b}$ ,  $t\bar{t}$  and  $Z$  + jets events. Their contribution is suppressed by requirements on lepton isolation and lepton impact parameter and their yield is estimated from control samples in data.

To help distinguish the Higgs signal from the dominant non-resonant  $ZZ^*$  background, both ATLAS [130] and CMS [128] use a matrix element likelihood approach to construct a kinematic discriminant built for each  $4\ell$  event based on the ratio of complete leading-order matrix elements  $|\mathcal{M}_{\text{sig}}|^2/|\mathcal{M}_{\text{bkg}}|^2$  for the signal ( $gg \rightarrow H \rightarrow 4\ell$ ) and background ( $q\bar{q} \rightarrow ZZ \rightarrow 4\ell$ ) hypotheses. The signal matrix element  $\mathcal{M}_{\text{sig}}$  is computed assuming  $m_H = m_{4\ell}$ . To further enhance the sensitivity to a signal, the ATLAS experiment uses the matrix element as an input variable to a Boosted Decision Tree, along with the transverse momentum and rapidity of the four-leptons system [130].

To enhance the sensitivity to VBF and VH production processes, the ATLAS and CMS experiments divide  $4\ell$  events into mutually exclusive categories. Events containing dijets with a large mass and pseudorapidity difference populate the VBF category. ATLAS requires the presence of an additional lepton in the VH category. In events with less than two jets, CMS uses the  $p_T^{\ell\ell}$  to distinguish between production via the gluon fusion and the VH/VBF processes.

Since the  $m_{4\ell}$  resolutions and the reducible background levels are different in the  $4\mu$ ,  $4e$  and  $2e2\mu$  subchannels, they are analyzed separately and the results are then combined.

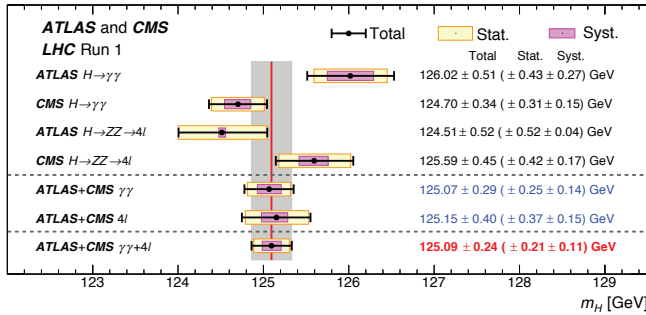
As shown in Fig. 11.3, the CMS experiment observes [128] its largest excess at  $m_H = 125.6$  GeV with an observed local significance of  $6.8\sigma$  to be compared with an expected significance of  $6.7\sigma$  at that mass. In their combined  $m_{4\ell}$  distribution, ATLAS observes [130] an excess at  $m_H = 125.36$  GeV with a local significance of  $8.1\sigma$ . The expected local significance for the SM Higgs boson at that mass is  $6.2\sigma$ . Both experiments also observe a clear peak at  $m_{4\ell} = 91$  GeV from  $Z/\gamma^*$  production at the expected SM rate [131].

The signal strength  $\mu$  for the inclusive  $H \rightarrow 4\ell$  production measured by the ATLAS and CMS experiments is  $1.44^{+0.40}_{-0.33}$  at  $m_H = 125.36$  GeV and  $0.93^{+0.29}_{-0.25}$  at  $m_H = 125.6$  GeV, respectively.

## III.2. Measurement of the Higgs boson mass

To measure the mass of the Higgs boson, ATLAS and CMS experiments rely on the two high mass resolution and sensitive channels,  $\gamma\gamma$  and  $ZZ$ . The approaches are very similar in these two analyses for both experiments, with subtle differences on

the use of categories, additional discriminating variables and per-event errors. These two channels are chosen for this precision measurement because they produce a narrow peak in mass with a resolution ranging from 1.4 GeV to 2 GeV for ATLAS and from 1.0 GeV to 2.8 GeV for CMS, where the best mass resolution is obtained for both experiments in the diphoton channel for central diphoton pairs (typically for events where both photons are not converted). For a model-independent mass measurement, the signal strengths in the  $\gamma\gamma$  and  $ZZ$  channels are assumed to be independent and not constrained to the expected rate ( $\mu = 1$ ) for the SM Higgs boson. The combined mass measured by ATLAS [132] and CMS [133] are  $125.36 \pm 0.37(\text{stat.}) \pm 0.18(\text{syst.})$  GeV and  $125.02^{+0.26}_{-0.27}(\text{stat.})^{+0.14}_{-0.15}(\text{syst.})$  GeV, respectively. In both experiments the measurements are dominated by the data statistics, however the systematic uncertainty is not negligible and is dominated by the precision in the knowledge of the photon energy or momentum scale. The ATLAS and CMS experiments have performed a combination of their mass measurements [134].



**Figure 11.4:** A compilation of the CMS and ATLAS mass measurements in the  $\gamma\gamma$  and  $ZZ$  channels, the combined result from each experiment and their combination. From Ref. [134]

Figure 11.4 summarizes these measurements and their combination [134]. The significance of the difference between the measurements of the masses in the  $\gamma\gamma$  and  $ZZ$  channels by the ATLAS experiment is  $1.97\sigma$  [132]. The ATLAS and CMS combined mass measurement:

$$m_H = 125.09 \pm 0.21(\text{stat.}) \pm 0.11(\text{syst.}) \text{ GeV}$$

reaches a precision of 0.2% and is dominated by statistical uncertainties.

In the diphoton channel, as is discussed in Section V.3.2 a mass shift is expected to be induced by the deformation of the mass lineshape of the signal in presence of background, from the interference between the Higgs boson production and the continuum irreducible background. It is a small but non negligible effect of approximately 35 MeV for a Higgs boson width close to that of the SM, but this effect could be larger if the width of the discovered particle were to be completely different. This effect estimated by ATLAS with a full simulation is still relatively small with respect to the total uncertainty on the mass and is therefore neglected.

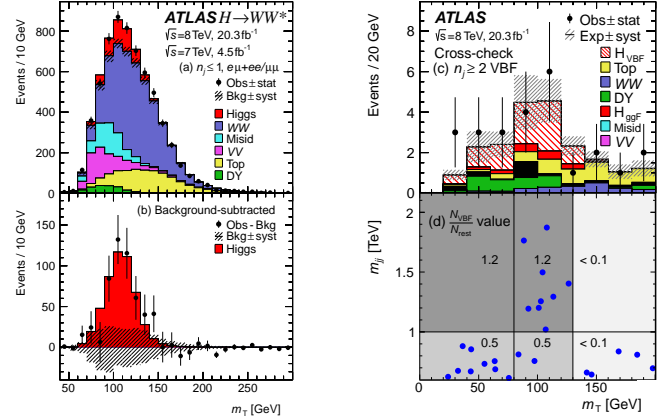
### III.3. $H \rightarrow W^+W^- \rightarrow \ell^+\nu\ell^-\bar{\nu}$

While the production rate in the  $H \rightarrow W^+W^- \rightarrow \ell^+\nu\ell^-\bar{\nu}$  channel is large, due to the presence of two neutrinos in the decay, the  $m_H$  resolution is quite poor ( $\approx 20\% m_H$ ) so the search requires fitting in several characteristic kinematic variables.

Experiments search for an excess of events with two leptons of opposite charge accompanied by missing energy and up to two jets. Events are divided into several categories depending on the lepton flavor combination ( $e^+e^-$ ,  $\mu^+\mu^-$  and  $e^\pm\mu^\mp$ ) and the number of accompanying jets ( $N_{\text{jet}} = 0, 1, \geq 2$ ). The  $N_{\text{jet}} \geq 2$  category is optimized for the VBF production process by selecting two leading

jets with a large pseudorapidity difference and with a large mass ( $m_{jj} > 500$  GeV).

Backgrounds contributing to this channel are numerous and depend on the category of selected events. Reducing them and accurately estimating the remainder is a major challenge in this analysis. For events with opposite-flavor lepton and no accompanying high  $p_T$  jets, the dominant background stems from non-resonant  $WW$  production. Events with same-flavor leptons suffer from large Drell-Yan contamination. The  $t\bar{t}$ ,  $Wt$  and  $W$  + jets (with the jet misidentified as a lepton) events contaminate all categories. Non-resonant  $WZ$ ,  $ZZ$  and  $W\gamma$  processes also contribute to the background at a sub-leading level.



**Figure 11.5:** (a) The  $m_T$  distribution for selected events summed over all lepton flavors and with  $\leq 1$  associated jets. (b) The residual of the data over the estimated SM background and the expectation from a SM Higgs boson with  $m_H = 125$  GeV indicating a clear excess with an event yield consistent with that from a SM Higgs boson [135]. The (c)  $m_T$  and (d)  $m_{jj}$  versus  $m_T$  distributions for the  $N_{\text{jet}} \geq 2$  VBF-enriched category for the 8 TeV data analysis. For each region in (d), the ratio  $N_{\text{VBF}}/N_{\text{rest}}$  is given ( $N_{\text{rest}}$  includes all production processes other than the VBF).

A requirement of large missing transverse energy ( $E_T^{\text{miss}}$ ) is used to reduce the Drell-Yan and multijet backgrounds. In the  $e^+e^-$  and  $\mu^+\mu^-$  categories, events with  $m_{\ell\ell}$  consistent with the  $Z$  mass are vetoed. The  $t\bar{t}$  background is suppressed by a veto against identified b-jets or low  $p_T$  muons (assumed to be coming from semileptonic b-hadron decays within jets) and tight isolation requirements diminish the  $W$ +jets background. The scalar nature of the Higgs boson and the  $V - A$  nature of the  $W$  boson decay implies that the two charged leptons in the final state are preferentially emitted at small angles with respect to each other. Therefore the dilepton invariant mass ( $m_{\ell\ell}$ ) and the azimuthal angle difference between the leptons ( $\Delta\phi_{\ell\ell}$ ) are used to discriminate between the signal and non-resonant  $WW$  events. The transverse mass, constructed from the dilepton  $p_T$  ( $p_T^{\ell\ell}$ ),  $E_T^{\text{miss}}$  and the azimuthal angle between  $E_T^{\text{miss}}$  and  $p_T^{\ell\ell}$ , is defined as  $m_T = \sqrt{2p_T^{\ell\ell}E_T^{\text{miss}}(1 - \cos\Delta\phi_{E_T^{\text{miss}}p_T^{\ell\ell}})}$  and serves as an effective discriminant against backgrounds. The transverse mass variable also tracks the Higgs boson mass but with a poor mass resolution. All residual background rates except for the small contributions from non-resonant  $WZ$ ,  $ZZ$  and  $W\gamma$  are evaluated from control samples devised from data.

The  $m_T$  distributions of the selected events in Run 1 data are shown in Fig. 11.5 for the ATLAS experiment. The 0-jet category is dominated by non-resonant  $WW$  background while  $t\bar{t}$  dominates the 1 and 2 jet categories. A clear excess over background expectation in the 0 and 1 jet categories is observed. An excess is also observed in the VBF-enriched 2-jets category. The observed event yield is consistent with the expectation from a 125 GeV SM Higgs boson.

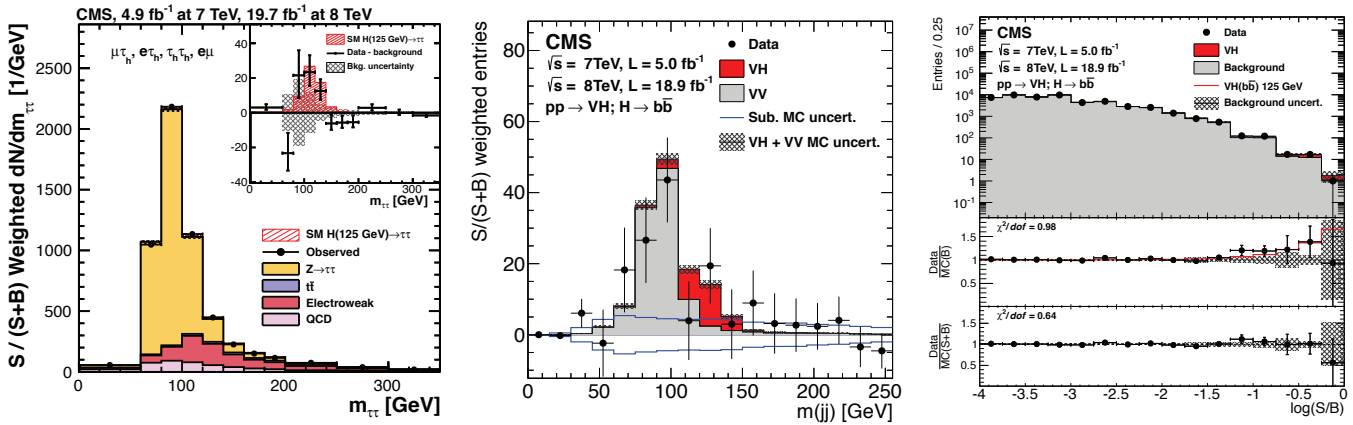
ATLAS fits the  $m_T$  distributions and observes [135] an excess at  $m_H = 125.36$  GeV with a local significance of  $6.1\sigma$  similar to that

expected from a 125 GeV SM Higgs boson. The measured inclusive signal strength is  $\mu = 1.09^{+0.23}_{-0.21}$ . In the VBF category an excess with a significance of  $3.2\sigma$  corresponding to a signal strength of  $\mu = 1.27^{+0.53}_{-0.45}$  is observed [135]. The CMS analysis of 0 and 1 jet categories, using all lepton flavor combinations, shows [136] an excess with an observed significance of  $4.3\sigma$ , lower than the expected sensitivity of  $5.8\sigma$  for a 125.6 GeV SM Higgs boson. CMS observes [136] no significant excess in the VBF production mode and sets a 95% CL limit on the signal strength of  $\mu_{\text{VBF}} < 1.7$  for  $m_H = 125.6$  GeV.

The ATLAS and CMS experiments have also searched for the associated Higgs boson production (VH) in this channel. The signal consists of up to three (WH) or four (ZH) high  $p_T$  isolated leptons with missing transverse energy and low hadronic activity. The major backgrounds stem from triboson and diboson production where each boson decays leptonically. ATLAS observes [137] an excess at  $m_H = 125.36$  GeV with a local significance of  $2.5\sigma$  corresponding to a  $\mu_{\text{VH}} = 3.0^{+1.6}_{-1.0}$ . CMS instead sets [136] a 95% CL limit of  $\mu_{\text{VH}} < 4.7$ .

resolution, which improves with the boost of the Higgs boson. The non-VBF categories are further subdivided according to the observed boost of the  $\tau^+\tau^-$  system. The 0-jet category which has the poorest signal/background ratio is used to constrain the background yields, the reconstruction efficiencies, and the energy scales. CMS primarily uses the reconstructed  $m_{\tau\tau}$  as the final discriminating variable [138] while the ATLAS experiment combines various kinematic properties of each event categories with multivariate techniques to build the final discriminant [139].

Searches for  $H \rightarrow \tau^+\tau^-$  decays in the VH production mode are performed in final states where the  $W$  or  $Z$  boson decays into leptons or jets. The irreducible background in this search arises from non-resonant  $WZ$  and  $ZZ$  diboson production. The reducible backgrounds originate from  $W$ ,  $Z$ , and  $t\bar{t}$  events that contain at least one fake lepton in the final state due to a misidentified jet. The shape and yield of the major backgrounds in each category is estimated from control samples in data. Contributions from non-resonant  $WZ$



**Figure 11.6:** (Left) The observed and predicted  $m_{\tau\tau}$  distributions for all  $H \rightarrow \tau^+\tau^-$  subchannels combined by the CMS experiment. The inset shows the difference between the observed data and the expected SM background contributions, together with the expected signal distribution for a SM Higgs boson with  $m_H = 125$  GeV [138]. (Center) The  $m_{b\bar{b}}$  distribution for the  $pp \rightarrow V(H \rightarrow b\bar{b})$  channels with all backgrounds except dibosons subtracted. The solid histograms for the backgrounds and the signal are summed cumulatively [140]. (Right) The combination of all  $pp \rightarrow V(H \rightarrow b\bar{b})$  channels into a single multivariate distribution. The two bottom panels show the ratio of the data to the background-only prediction (above) and to the predicted sum of background and SM Higgs boson signal with a mass of 125 GeV (below).

### III.4. Decays to fermions

At hadron colliders, the most promising channel for probing the coupling of the Higgs field to the quarks and leptons are  $H \rightarrow b\bar{b}$  and  $H \rightarrow \tau^+\tau^-$ , respectively. For a Higgs boson with  $m_H \approx 125$  GeV, the branching fraction to  $b\bar{b}$  is about 57% and to  $\tau^+\tau^-$  is about 6%. Nevertheless, the presence of very large backgrounds makes the isolation of a Higgs boson signal in these channels quite challenging.

#### III.4.1. $H \rightarrow \tau^+\tau^-$

In the  $H \rightarrow \tau\tau$  search,  $\tau$  leptons decaying to electrons ( $\tau_e$ ), muons ( $\tau_\mu$ ) and hadrons ( $\tau_{had}$ ) are considered. The  $\tau^+\tau^-$  invariant mass ( $m_{\tau\tau}$ ) is reconstructed from a kinematic fit of the visible products from the two  $\tau$  leptons and the missing energy observed in the event. Due to the presence of missing neutrinos, the  $m_{\tau^+\tau^-}$  resolution is poor ( $\approx 15\%$ ). As a result, a broad excess over the expected background in the  $m_{\tau\tau}$  distribution is searched for. The major sources of background stem from Drell-Yan  $Z \rightarrow \tau^+\tau^-$  and  $Z \rightarrow e^+e^-$ ,  $W$ +jets,  $t\bar{t}$  and multijet production. Events in all subchannels are divided into categories based on the number and kinematic properties of additional energetic jets in the event. The sensitivity of the search is generally higher for categories with one or more additional jets. The VBF category, consisting of a  $\tau$  pair with two energetic jets separated by a large pseudorapidity, has the best signal-to-background ratio and search sensitivity, followed by the  $\tau^+\tau^-$ +1 jet category. The signal to background discrimination relies in part on the  $m_{\tau\tau}$

and  $ZZ$  diboson production are estimated from simulations but corrected for reconstruction efficiency using control samples formed from observed data.

Figure 11.6 shows the CMS [138]  $m_{\tau\tau}$  distributions combining all categories, weighing the distributions in each category of each subchannel by the ratio between the expected signal and background yields for that category. The inset plot shows the difference between the observed and expected background distributions, together with the expected distribution for a SM Higgs boson signal with  $m_H = 125$  GeV. The significance of the observed excess at  $m_H = 125$  GeV is 3.4 standard deviations, close to the expected sensitivity, and corresponds to a signal strength of  $\mu = 0.86 \pm 0.29$ . At  $m_H = 125.36$  GeV, the observed (expected) deviation from the background-only hypothesis at ATLAS corresponds to a local significance of 4.5 (3.4) standard deviations and the best fit value of the signal strength is  $\mu = 1.43^{+0.43}_{-0.37}$  [139].

When the ATLAS and CMS  $H \rightarrow \tau\tau$  measurements are combined [141], the significance of the observed excess corresponding to  $m_H = 125.09$  GeV is 5.5 standard deviations and the combined signal strength is  $\mu = 1.11^{+0.24}_{-0.22}$ .

#### III.4.2. $H \rightarrow b\bar{b}$

The production mode  $gg \rightarrow H$  with  $H \rightarrow b\bar{b}$  is overwhelmed by the background from the inclusive production of  $p\bar{p} \rightarrow b\bar{b} + X$  via the strong interaction. The associated production modes WH and



ZH (collectively termed VH modes) allow use of the leptonic  $W$  and  $Z$  decays for triggering, and to purify the signal and reject QCD backgrounds. The  $W$  bosons are reconstructed via their leptonic decay  $W \rightarrow \ell \bar{\nu}_\ell$  where  $\ell = e, \mu$  or  $\tau$ . The  $Z$  bosons are reconstructed via their decay into  $e^+e^-$ ,  $\mu^+\mu^-$  or  $\nu\bar{\nu}$ . The Higgs boson candidate mass is reconstructed from two b-tagged jets in the event. Backgrounds arise from production of  $W$  and  $Z$  bosons in association with gluon, light and heavy-flavored jets (V+jets),  $t\bar{t}$ , diboson ( $ZZ$  and  $WZ$  with  $Z \rightarrow b\bar{b}$ ) and QCD multijet processes. Due to the limited  $m_{b\bar{b}}$  mass resolution, a SM Higgs boson signal is expected to appear as a broad enhancement in the reconstructed dijet mass distribution. The crucial elements in this search are b-jet tagging with high efficiency and low fake rate, accurate estimate of b-jet momentum and estimate of backgrounds from various signal depleted control samples constructed from data.

At the Tevatron, the  $H \rightarrow b\bar{b}$  channel contributes the majority of the Higgs boson search sensitivity below  $m_H = 130$  GeV. The CDF and D0 experiments use multivariate analysis (MVA) techniques that combine several discriminating variables into a single final discriminant used to separate signal from background. Each channel is divided into exclusive subchannels according to various lepton, jet multiplicity, and b-tagging characteristics in order to group events with similar signal-to-background ratio and thus optimize the overall search sensitivity. The combined CDF and D0 data show [142, 126] an excess of events with respect to the predicted background in the 115–140 GeV mass range in the most sensitive bins of the discriminant distributions suggesting the presence of a signal. At  $m_H = 125$  GeV the local significance of the excess is 3.0 standard deviations. At that mass, the observed signal strength  $\mu = 1.59^{+0.69}_{-0.72}$ .

To reduce the dominant V+jets background, following Ref. [143], the LHC experiments select a region in VH production phase space where the vector boson is significantly boosted and recoils from the  $H \rightarrow b\bar{b}$  candidate with a large azimuthal angle  $\Delta\phi_{VH}$ . For each channel, events are categorized into different  $p_T(V)$  regions with varying signal/background ratios. Events with higher  $p_T(V)$  have smaller backgrounds and better  $m_{b\bar{b}}$  resolution. CMS uses [140] MVA classifiers based on kinematic, topological and quality of b-jet tagging and trained on different values of  $m_H$  to separate Higgs boson signal in each category from backgrounds. The MVA outputs for all categories are then fit simultaneously. Figure 11.6(right) shows the combined MVA output of all categories where events are gathered in bins of similar expected signal-to-background ratios as predicted by the MVA discriminants. The excess of events observed in bins with the largest signal-to-background ratios is consistent with the production of a 125 GeV SM Higgs boson with a significance of 2.1 standard deviations. The observed signal strength at 125 GeV is  $\mu = 1.0 \pm 0.5$ . Figure 11.6(center) shows the  $m_{b\bar{b}}$  distribution for all categories combined, weighted by the signal-to-background ratio in each category, with all backgrounds except dibosons subtracted. The data show the clear presence of a diboson ( $W/Z + Z \rightarrow b\bar{b}$ ) signal, with a rate consistent with the SM expectation, together with an excess that agrees with that expected from the production of a 125 GeV SM Higgs boson. The nominal results from ATLAS are also based on a combination [144] of (i) a multivariate analysis in addition to  $m_{b\bar{b}}$  and b-tagging information and (ii) a statistical analysis of their 7 TeV data centered on  $m_{b\bar{b}}$  as the main discriminant. In both cases customized control samples devised from data are used to constrain the contributions of the dominant background processes. The net observed(expected) deviation from background-only hypothesis corresponds to a significance of 1.4(2.6) standard deviations and a signal strength of  $\mu = 0.5 \pm 0.4$ .

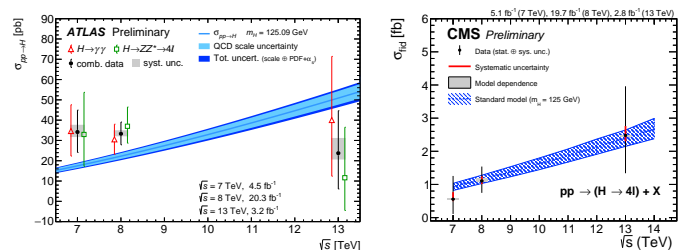
In their 8 TeV data, CMS has also searched for  $H \rightarrow b\bar{b}$  in the VBF production mode [145]. The event topology consists of two “VBF-tagging” energetic light-quark jets in the forward and backward direction relative to the beam direction and two b-tagged jets in the central region of the detector. Due to the electroweak nature of the process, for the signal events, no energetic jet activity is expected in the rapidity gap between the two “VBF-tagging” jets. The dominant background in this search stems from QCD production of multijet events and the hadronic decays of vector bosons accompanied by

additional jets. A contribution of Higgs boson events produced in the ggF process but with two or more associated jets is expected in the signal sample. The signal is expected as a broad enhancement in the  $m_{b\bar{b}}$  distribution over the smoothly falling contribution from the SM background processes. The observed (expected) excess corresponding to  $m_H = 125$  GeV was 2.2 (0.8) standard deviations corresponding to a signal strength of  $\mu = 2.8^{+1.6}_{-1.4}$ . Combining with the result of the CMS VH analysis yields a signal strength signal  $\mu = 1.0 \pm 0.4$  and the local significance of the excess improves marginally to 2.6 standard deviations.

### III.5. First results on the main production and decay channels at 13 TeV

After a period of long shutdown between 2013 and 2015 devoted to the consolidation of the machine, in Spring 2015 the LHC delivered  $pp$  collisions at an unprecedented centre-of-mass energy of 13 TeV. During this period the ATLAS and CMS experiments have collected datasets corresponding to integrated luminosities of 2.3 to 3.2 fb<sup>-1</sup> for CMS and ATLAS, respectively. The first preliminary measurements of Higgs boson production at this increased centre-of-mass energy are arriving while this review is being finalized. Only a concise section and an update in the associated production with a top-quark pair are therefore devoted to these results.

The two high-resolution channels  $H \rightarrow \gamma\gamma$  and  $H \rightarrow ZZ^* \rightarrow \ell^+\ell^-\ell'^+\ell'^-$  have been measured both by the ATLAS and CMS experiments. With the increase in production cross sections, a fair sensitivity is expected in these two channels even with this limited amount of data. For the  $H \rightarrow \gamma\gamma$  channel ATLAS has produced a fully inclusive analysis in order to measure a fiducial cross section with a sensitivity of 1.9 $\sigma$  and an observed excess of 1.8 $\sigma$  [146]. The measurement of the total cross section, as extrapolated from the fiducial region is shown in Fig. 11.7(left). CMS has produced an analysis with event classification that has reached a sensitivity of 2.7 $\sigma$  and has observed an overall excess of 1.7 $\sigma$  [147]. For the  $H \rightarrow ZZ^*$  channel, ATLAS has also produced a fully inclusive analysis with a sensitivity of 2.8 $\sigma$  and no significant excess with respect to the background has been observed [148]. This outcome is however compatible with the presence of a signal at the 1.4 $\sigma$  level. The corresponding measurement of the total cross section is illustrated in Fig. 11.7(left). The CMS analysis in this channel is also inclusive and uses additional kinematic discriminants to reach a sensitivity of 3.4 $\sigma$  [149]. CMS observes an excess with a significance of 2.5 $\sigma$ . In this channel, CMS also measures a fiducial cross section as shown in Fig. 11.7(right). The ATLAS experiment has also performed a combination of the total cross sections in these two channel at 7, 8 and 13 TeV [150]. The results of these combinations are shown in Fig. 11.7(left).



**Figure 11.7:** (Left) Total cross sections measured by ATLAS at 7, 8 and 13 TeV in the  $H \rightarrow \gamma\gamma$  and  $H \rightarrow ZZ^*$  channels and their combinations. (Right) The fiducial cross section measured in the  $H \rightarrow ZZ^*$  channel by CMS at 7, 8 and 13 TeV.

The CMS experiment has investigated two additional channels. The first is the  $H \rightarrow WW^* \rightarrow \ell\nu\ell\nu$  in categories of up to one jet, with a sensitivity of 2 $\sigma$  [151] and has observed only a very mild excess of 0.7 $\sigma$ . The second is the  $H \rightarrow b\bar{b}$  decay mode in the VBF production sensitive only to approximately twice the SM production rate. The observation in this channel is compatible both with the background and the background in presence of a SM signal [152].

### III.6. Higgs production in association with top quarks or in top decays

#### III.6.1. The associated production with top quark pairs

As discussed in Section II, the coupling of the Higgs particle to top quarks plays a special role in the electroweak breaking mechanism and in its possible extensions. Substantial indirect evidence of this coupling is provided by the compatibility of observed rates of the Higgs boson in the principal discovery channels, given that the main production process – the gluon fusion – is dominated by a top quark loop. Direct evidence of this coupling at the LHC and the future  $e^+e^-$  colliders will be mainly available through the  $t\bar{t}H$  final state and will permit a clean measurement of the top quark-Higgs boson Yukawa coupling. The  $t\bar{t}H$  production cross section at the LHC is tiny in comparison with the ggF or even  $VH$  production modes. The production cross section for a 125 GeV Higgs boson in  $pp$  collisions at  $\sqrt{s} = 8$  TeV of about 130 fb makes it challenging to measure the  $t\bar{t}H$  process with the LHC Run 1 dataset. It is thus imperative to target every accessible experimental signature. The analyses channels for such complex final states can be separated in four classes according to the decays of the Higgs boson. In each of these classes, most of the decay final states of the top quarks are considered. The topologies related to the decays of the top quarks are denoted 0L, 1L and 2L, for the fully hadronic, semi-leptonic and dilepton decay final states of the  $t\bar{t}$ , respectively.

The first analysis in this set is the search for  $t\bar{t}H$  production in the  $H \rightarrow \gamma\gamma$  channel. This analysis relies on the search for a narrow mass peak in the  $m_{\gamma\gamma}$  distribution. The background is estimated from the  $m_{\gamma\gamma}$  sidebands. The sensitivity in this channel is mostly limited by the available statistics. The second is the search in the  $H \rightarrow b\bar{b}$  channel. This search is extremely intricate due to the large backgrounds, both physical and combinatorial in resolving the  $b\bar{b}$  system related to the Higgs particle, in events with six jets and four  $b$ -tagged jets which are very hard to simulate. With the current dataset, the sensitivity of this analysis is severely impacted by the systematic uncertainties on the background predictions. The third channel is a specific search for  $\tau^+\tau^-$  where the two tau leptons decay to hadrons. Finally, the  $W^+W^-$ ,  $\tau^+\tau^-$  and  $ZZ$  final states can be searched for inclusively in multilepton event topologies. The corresponding  $t\bar{t}H$  modes can be decomposed in terms of the decays of the Higgs boson and those of the top quarks as having two  $b$ -quarks and four  $W$  bosons (or two  $W$  and two taus, or two  $W$  and two  $Z$ ) in the final state.

CMS combines these four sets of measurements [153] and reports a 95% CL upper limit on the signal strength value of  $\mu_{t\bar{t}H} < 4.5$ . ATLAS reports [155–157] 95% CL upper limits on the signal strengths of 6.7, 6.4 and 4.7 for  $H \rightarrow \gamma\gamma$ ,  $H \rightarrow b\bar{b}$  and  $H \rightarrow$  multilepton decay final states, respectively. The CMS experiment has also updated the  $t\bar{t}H(\rightarrow b\bar{b})$  analysis using the matrix element method, aiming at an optimal separation between the signal and the dominant  $t\bar{t}$  production in association with heavy flavor quarks in the final state [154].

#### III.6.2. The associated production with a single top quark

An additional production mode of the Higgs boson in association with a top quark is the single top associated production mode. There is an interesting similarity between this production mode and the  $H \rightarrow \gamma\gamma$  decay mode. Both processes proceed through either the top Yukawa coupling or the interaction of the Higgs boson with the  $W$ -boson, with a negative interference between the two. Representative Feynman diagrams for this production process are shown in Fig. 11.8. Contrary to the diphoton decay channel, in this production mode the interference occurs at the tree level. This process can be used to further discriminate a negative relative sign between the couplings of the Higgs boson to fermions and its couplings to gauge bosons.

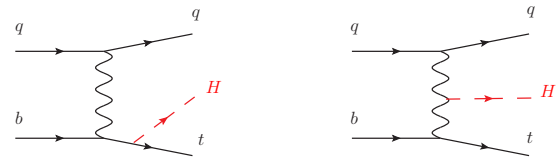
The ATLAS experiment has re-interpreted its Run 1 search of the diphoton decay channel in the  $t\bar{t}H$  production in terms of  $tH$  production [155] and has produced 95% CL upper limits on the Higgs boson production cross section with respect to the rates expected for a given sign of the top Yukawa coupling. The result is not strong enough to exclude, at 95% CL, a negative top Yukawa coupling with an absolute strength equal to that of the SM. The excluded range in

**Table 11.5:** Summary of the results of searches for a Higgs boson in association with a top quark pair by the ATLAS and CMS collaborations. The results are given in terms of measured signal strength. The results of subchannels including hadronically decaying taus, which are less sensitive, are not reported in this table but can be found in the corresponding references.

	ATLAS (7 and 8 TeV)	CMS (7 and 8 TeV)	CMS (13 TeV)
$t\bar{t}(H \rightarrow \gamma\gamma)$	$1.3^{+2.6}_{-1.7}^{+2.5}_{-1.7}$	$1.2^{+2.5}_{-1.7}^{+2.6}_{-1.8}$	—
$t\bar{t}(H \rightarrow b\bar{b})$ -0L	$1.6 \pm 0.8 \pm 2.5$	—	—
$t\bar{t}(H \rightarrow b\bar{b})$ -1L	$1.2 \pm 0.8 \pm 0.8$	$1.7^{+2.0}_{-1.8}$	$-0.4^{+2.1}_{-2.1}$
$t\bar{t}(H \rightarrow b\bar{b})$ -2L	$2.8; \pm 1.4 \pm 2.0$	$1.0^{+3.3}_{-3.0}$	$-4.7^{+3.7}_{-3.8}$
$t\bar{t}(H \rightarrow b\bar{b})$	$1.4 \pm 1.0 \pm 0.6$	$1.6^{+1.6}_{-1.5}$	$-2.0 \pm 1.8$
$t\bar{t}(H \rightarrow 4\ell)$	$2.8^{+2.0}_{-1.7}^{+0.9}_{-0.6}$	$-4.7^{+5.0}_{-1.3}$	—
$t\bar{t}(H \rightarrow 3\ell)$	$2.8^{+2.0}_{-1.7}^{+0.9}_{-0.6}$	$3.1^{+2.4}_{-2.0}$	$5.8^{+3.3}_{-2.7}$
$t\bar{t}(H \rightarrow SS2\ell)$	$2.8^{+1.5}_{-1.4}^{+1.5}_{-1.3}$	$5.3^{+2.1}_{-1.8}$	$-0.5^{+1.0}_{-0.7}$
$t\bar{t}(H \rightarrow WW/\tau\tau/ZZ)$	$1.4 \pm 0.6 \pm 1.0$	—	$0.6^{+1.4}_{-1.1}$
Combination	$1.7 \pm 0.5 \pm 0.8$	$2.8^{+1.0}_{-0.9}$	—

the ratio of the Higgs-top Yukawa coupling to that of the SM one,  $\kappa_t$ , at the 95% CL, is  $]-\infty, -1.3] \cup [8, \infty[$ .

The CMS experiment has produced a search with the Run 1 data exploiting a variety of Higgs boson decay modes resulting in final states with photons, bottom quarks, and multiple charged leptons, including tau leptons. The analysis is optimized for the opposite sign of the top Yukawa coupling with respect to that in the SM, and corresponding to a large enhancement of the signal cross section. The expected sensitivity of this analysis in terms of exclusion of the ratio of the cross section to the expected SM cross section is 2 for  $\kappa_t = -1$ , while the observed exclusion limit is 2.8.



**Figure 11.8:** Feynman diagrams contributing to the Higgs production in association with a single top-quark through the top Yukawa coupling (left) and through the Higgs coupling to the  $W$  gauge boson (right).

#### III.6.3. Flavor changing neutral current decays of the top quark

The discovery of the Higgs boson at a mass smaller than the top quark mass opened a new decay channel for the top quark. The decays of the top quark to a Higgs boson and a charm or an up quark proceed through a Flavor Changing Neutral Current (FCNC) which are forbidden at the tree level and suppressed at higher orders through the Glashow–Iliopoulos–Maiani (GIM) mechanism [3]. The SM prediction for these branching fractions is  $\text{BR}(t \rightarrow Hc) = 10^{-15}$  and two orders of magnitude less for the  $Hu$  final state. These decay channels of the top quark are, however, very interesting to probe possible FCNC interactions in the Higgs Yukawa couplings to the quark sector.

The ATLAS experiment has searched for FCNC top decays specifically in channels involving a Higgs boson with subsequent decays to two photons [158] and a pair of  $b$ -quarks [159]. It has also reinterpreted a search for the  $t\bar{t}H$  production in the multilepton final state (discussed in Section III.6.1) [157]. The latter channel covers Higgs boson decays to a pair of  $W$ -bosons and a pair of

taus. No significant excess was observed in any of the specific channels (as discussed in Section III.6.1, a slight excess is observed in the  $ttH$  multilepton channel) and 95% CL upper limits are set on  $\text{BR}(t \rightarrow Hc) < 0.46\%$  with an expected sensitivity of 0.25% and  $\text{BR}(t \rightarrow Hu) < 0.45\%$  with an expected sensitivity of 0.29%. The CMS experiment has performed a search for these FCNC top decays in the diphoton and multilepton channels [160], yielding a 95% CL upper limit on  $\text{BR}(t \rightarrow Hc) < 0.56\%$  with an expected sensitivity of 0.65%.

From these limits on branching fractions, constraints on non flavor-diagonal Yukawa couplings of a FCNC sector Lagrangian of the form:

$$\mathcal{L}_{\text{FCNC}} = \lambda_{tch} \bar{t} H c + \lambda_{tuh} \bar{t} H u + h.c.$$

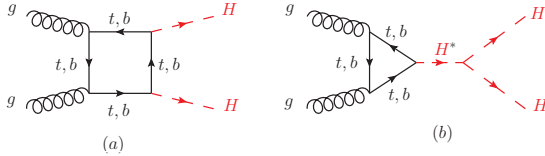
can be derived. The 95% CL observed (expected) upper limits from ATLAS on the  $|\lambda_{tch}|$  and  $|\lambda_{tuh}|$  couplings are 0.13 (0.10) and 0.13 (0.10), respectively.

### III.7. Searches for other rare production modes

#### III.7.1. Searches for Higgs boson pair production

Higgs boson pair production in the SM is rare. It is however a very interesting final state to search in two specific modes: (i) the search for non-resonant production of the Higgs boson pair and (ii) the search for resonant production of two Higgs bosons in the decay of a heavier particle.

Non-resonant Higgs pair production is an interesting milestone in the study of prospects for constraining Higgs self-couplings. In the SM the main non-resonant production mode of two Higgs bosons in the final state proceeds through a loop (mainly of top quarks) (Fig. 11.9a). Another production mode is via the trilinear coupling of the Higgs boson (Fig. 11.9b), whose amplitude is not negligible compared to the former. These diagrams interfere negatively making the overall production rate smaller than what would be expected in the absence of a trilinear coupling. The sensitivity to the trilinear coupling will be discussed in Section III.7.1.ii.



**Figure 11.9:** Feynman diagrams contributing to Higgs boson pair production through (a) a top- and b-quark loop and (b) through the self couplings of the Higgs boson.

#### (i) Searches for Higgs boson pair production

The searches for Higgs boson pair production both resonant and non-resonant are very interesting probes for a variety of theories beyond the SM, and can be done in a large number of Higgs boson decay channels. The ATLAS collaboration has searched both for resonant and non resonant Higgs boson pair production in the following channels: (i)  $HH \rightarrow b\bar{b}\gamma\gamma$  [161]; (ii)  $HH \rightarrow b\bar{b}\tau^+\tau^-$  [162]; (iii)  $HH \rightarrow b\bar{b}b\bar{b}$  [163]; and (iv)  $HH \rightarrow WW^*\gamma\gamma$  [162]. The CMS collaboration has only performed searches for resonant Higgs boson pair production in a large variety of decay modes: (i) in final states containing multiple leptons (electrons or muons) covering the  $WW^*WW^*$ ,  $WW^*ZZ^*$ ,  $ZZ^*ZZ^*$ ,  $ZZ^*\tau^+\tau^-$ ,  $WW^*\tau^+\tau^-$ ,  $ZZ^*b\bar{b}$ ,  $\tau^+\tau^-\tau^+\tau^-$  channels [164]; (ii) in final states with a di-photon pair compatible with being produced in the decay of the Higgs boson and one lepton covering the  $\gamma\gamma WW^*$ ,  $\gamma\gamma ZZ^*$ ,  $\gamma\gamma\tau^+\tau^-$  channels [164]; (iii) in the  $b\bar{b}\tau^+\tau^-$  channel [165]; and (iv) in the  $b\bar{b}b\bar{b}$  channel [166]. A summary of the channels searched for at the LHC is given in Table 11.6. The results and interpretation of the search for resonant Higgs boson pair production are discussed in Section VII.8.i.d.

**Table 11.6:** Summary of the final states investigated in the search for Higgs boson pair production by ATLAS (A) and CMS (C).

	$b\bar{b}$	$\tau^+\tau^-$	$4\ell$	$\ell^+\nu\ell^-\bar{\nu}$	$\gamma\gamma$
$b\bar{b}$	A [163], C [166]	A [162], C [165]	C [164]	C [164]	A [161]
$\tau^+\tau^-$	–	C [164]	C [164]	C [164]	C [164]
$4\ell$	–	–	C [164]	C [164]	C [164]
$\ell^+\nu\ell^-\bar{\nu}$	–	–	–	C [164]	A [162], C [164]

#### (ii) Measuring Higgs self couplings

The Higgs boson self coupling is an extremely important direct probe of the Higgs potential. The measurement of the quartic coupling in  $HHH$  final states is essentially deemed to be impossible at the HL-LHC. The possibility of measuring the trilinear coupling of the Higgs boson at the LHC in  $HH$  final states is studied in detail.

In the SM the Higgs boson pair production through the trilinear Higgs has an on-shell component and a large off-shell component. The on-shell  $H \rightarrow H^*H^*$  is strongly disfavored, requiring two off-shell Higgs bosons in the final state. The sensitivity region to the trilinear coupling production as in Fig. 11.9-b, is mainly in the kinematic region where the two Higgs boson in the final state are on-shell and the Higgs boson acts as a propagator (off-shell). As discussed in the introduction to this section, this process interferes negatively with the background Higgs boson pair production (Fig. 11.9a). In the SM hypothesis sensitivity to the trilinear coupling requires the measurement of a deficit in the Higgs boson pair production, in a similar way as the off-shell couplings measurement as explained in Section V.2. Given the current sensitivity of Higgs boson pair production measurements discussed in the previous Section III.7.1.i, only projections for the high Luminosity LHC are considered with an integrated luminosity of  $3\text{ab}^{-1}$  and therefore in high pile-up conditions. Three channels have been investigated: (i) the  $HH \rightarrow b\bar{b}\gamma\gamma$ ; (ii) the  $HH \rightarrow b\bar{b}\tau^+\tau^-$ ; and (iii) the  $HH \rightarrow b\bar{b}W^+W^-$ . The prospects in channel (i) have been studied by both the ATLAS [167] and the CMS [168] collaborations, yielding a sensitivity of  $1.3\sigma$  and  $1.6\sigma$  respectively to overall Higgs boson pair production. The ATLAS and CMS collaborations have studied the channel (ii) yielding a sensitivity of  $0.6\sigma$  [169] and  $0.9\sigma$  [168] to Higgs boson pair production respectively. Only the CMS collaboration has studied the channel (iii) showing its low sensitivity [168]. It should be noted that there is a large uncertainty on these projections related both to the modeling of signal and the backgrounds, the very difficult high pile-up environment (both for reconstruction and trigger) and the design of the upgraded detectors. As discussed in Section III.7.1.i, more channels are possible and deserve to be studied in detail.

The measurements of the trilinear coupling requires to separate the contributions to the overall Higgs boson pair production and in particular measure the deficit expected in the case of SM couplings. The ATLAS collaboration has estimated the sensitivity to the trilinear  $\lambda_{HHH}$  coupling to exclusion regions of  $\lambda_{HHH}/\lambda_{HHH}^{\text{SM}}$  at 95% CL of  $]-\infty, -1.3] \cup [8.6, \infty[$  with the  $b\bar{b}\gamma\gamma$  channel [167] and  $]-\infty, -1.4] \cup [12.0, \infty[$  with the  $b\bar{b}\tau^+\tau^-$  channel [169]. Measuring the Higgs boson trilinear coupling will be very difficult at the High Luminosity LHC.

### III.8. Searches for rare decays of the Higgs boson

#### III.8.1. $H \rightarrow Z\gamma$

The search for  $H \rightarrow Z\gamma$  is performed in the final states where the  $Z$  boson decays into opposite sign and same flavor leptons ( $\ell^+\ell^-$ ),  $\ell$  here refers to  $e$  or  $\mu$ . While the branching fraction for  $H \rightarrow Z\gamma$  is comparable to  $H \rightarrow \gamma\gamma$  (about  $10^{-3}$ ) at  $m_H = 125\text{GeV}$ , the observable signal yield is brought down by the small branching ratio of  $Z \rightarrow (e^+e^- + \mu^+\mu^-) = 6.7 \times 10^{-2}$ . In these channels, the  $m_{\ell\ell\gamma}$  mass resolution is excellent (1-3%) so the analyses search for a narrow mass peak over a continuous background. The major backgrounds arise from the  $Z + \gamma$  final state radiation in Drell-Yan decays and  $Z + \text{jets}$  processes where a jet is misidentified as a photon. The ratio of signal over background in this channel is typically of the order of 0.5%. In



a narrow window of a few GeV around 125 GeV, several hundreds of events are expected.

Events are divided into mutually exclusive categories on the basis of the expected  $m_{Z\gamma}$  resolution and the signal-to-background ratio. A VBF category is formed for  $H \rightarrow Z\gamma$  candidates which are accompanied by two energetic jets separated by a large pseudorapidity. While this category contains only about 2% of the total event count, the signal-to-noise is about an order of magnitude higher. The search for a Higgs boson is conducted independently in each category and the results from all categories are then combined.

No excess of events is observed in either ATLAS or CMS experiments. The CMS expected and observed 95% CL upper limits for  $m_H = 125$  GeV [170] on the signal strength  $\mu$  are 10 and 9.5 respectively. The ATLAS expected and observed upper limits [171] on the signal strength  $\mu$  are 9 and 11 respectively for a SM  $m_H = 125.5$  GeV.

### III.8.2. $H \rightarrow \mu^+\mu^-$

$H \rightarrow \mu^+\mu^-$  is the only channel where the Higgs coupling to second generation fermions can be measured at the LHC. The branching fraction in this channel for a 125 GeV SM Higgs boson is  $2.2 \times 10^{-4}$ , about ten times smaller than that for  $H \rightarrow \gamma\gamma$ . The dominant and irreducible background arises from the  $Z/\gamma^* \rightarrow \mu^+\mu^-$  process which has a rate several orders of magnitude larger than that from the SM Higgs boson signal. Due to the precise muon momentum measurement achieved by ATLAS and CMS, the  $m_{\mu^+\mu^-}$  mass resolution is excellent ( $\approx 2-3\%$ ). A search is performed for a narrow peak over a large but smoothly falling background. For optimal search sensitivity, events are divided into several categories. To take advantage of the superior muon momentum measurement in the central region, the two experiments subdivide events by the pseudorapidity of the muons. To suppress the Drell-Yan background, ATLAS requires  $p_T^{\mu^+\mu^-} > 15$  GeV while CMS separates them into two  $p_T^{\mu^+\mu^-}$  based categories. CMS further categorizes events by the number and the topology of additional energetic jets in the event.

No excess in the  $m_{\mu^+\mu^-}$  spectrum is observed near 125 GeV. From an analysis of their Run 1 data, ATLAS sets [172] an observed (expected) 95% CL upper limit on the signal strength  $\mu < 7.0$  (7.2). The CMS analysis [173] of their 7 and 8 TeV data sets an observed (expected) limit of  $\mu < 7.4$  (6.5).

### III.8.3. $H \rightarrow e^+e^-$

A search similar to the  $H \rightarrow \mu^+\mu^-$ , is performed by CMS in the di-electron channel [173]. In this search channel there the contribution from the peaking background from Higgs boson decays to diphoton mis-identified as di-electrons (when mostly converted photons are faking electrons) needs to be assessed. The sensitivity to the SM Higgs decays is negligible given the extremely small branching fraction to  $e^+e^-$ , approximately 40,000 times smaller than the branching fraction to dimuons. It is nevertheless interesting to probe this decay channel to search for potential large anomalous couplings. Assuming a SM Higgs boson production cross section, the observed limit on the branching fraction at the 95% CL is 0.0019 [173], five orders of magnitude larger than the expected SM prediction.

### III.8.4. Lepton flavor violating (LFV) Higgs boson decays

Given the Yukawa suppression of the couplings of the Higgs boson to quarks and leptons of the first two generations and the small total width of the Higgs boson, new physics contributions could easily have sizable branching fractions. One very interesting possibility is the Lepton Flavor Violating (LFV) decays of the Higgs boson, in particular in the  $\tau\mu$  and  $\tau e$  modes. These decays are suppressed in the SM but could be enhanced in theories such as two-Higgs-doublet models (discussed in Section VII).

There are already fairly strong constraints on LFV Yukawa couplings  $|Y_{\tau\mu}|$  from channels such as the  $\tau \rightarrow 3\mu$  or  $\tau \rightarrow \mu\gamma$ , or a re-interpretation of the search for Higgs decays to  $\tau^+\tau^-$ . A direct search at the LHC however complements these indirect limits. The search for LFV decays in the  $\tau\mu$  channel have been done with the Run 1 dataset in several channels according to the subsequent decay of the

	ATLAS $\tau_{\text{had}}\mu$	ATLAS $\tau_{\text{lep}}\mu$	ATLAS $\tau\mu$	CMS $\tau\mu$
BR( $H \rightarrow \tau\mu$ )	$(0.77 \pm 0.62)\%$	$(0.03^{+0.88}_{-0.86})\%$	$(0.53 \pm 0.51)\%$	$(0.84^{+0.39}_{-0.37})\%$
95% CL Expected	1.24%	1.73%	1.01%	0.75%
95% CL Observed	1.85%	1.79%	1.43%	1.51%

**Table 11.7:** Summary of the results of searches for lepton flavor violating decays of the Higgs boson in the  $\tau\mu$  channel from ATLAS and CMS.

$\tau$ . The results from CMS [174] and for ATLAS for the hadronic [175], the leptonic [176] decays of the tau, and their combination [176] are reported in Table 11.7. It is interesting to note that the analysis strategies for the di-lepton  $\tau_{\text{lep}}\mu$  channel are very different between the ATLAS [176] and CMS [174].

As shown in Table 11.7 a small excess in this channel is observed by CMS with a significance of  $2.5\sigma$ , while in ATLAS the excess is smaller and of the order of  $1\sigma$ .

The ATLAS collaboration has also reported results on the search for the LFV Higgs boson decays in the  $\tau e$  channel [176], yielding an observed (expected) limit of 1.04% (1.21%).

### III.8.5. Probing charm- and light-quark Yukawa couplings

Probing the Yukawa couplings to quarks of the second or even the first generation is extremely challenging given the overwhelming background and the much smaller signal rates. The possibility of probing the Yukawa coupling to the charm has been discussed in [177] where indirect bounds on the charm Yukawa coupling are estimated from a combined fit to the Higgs data. The direct impact of Higgs decays to a pair of charm quarks on the direct search for  $H \rightarrow b\bar{b}$  is also investigated.

Another possibility to access the Higgs Yukawa coupling has been discussed in [178], through the decays of the Higgs boson to a final state with charmonium:  $H \rightarrow J/\psi\gamma$ . Higgs decays in this final state have been searched for by the ATLAS collaboration [179]. The sensitivity of this analysis is however several orders of magnitude above the branching fraction estimated for the SM coupling  $\text{BR}(H \rightarrow J/\psi\gamma) = (2.8 \pm 0.2) \times 10^{-6}$  [178]. The ATLAS collaboration [179] has also searched for Higgs decays to  $\Upsilon(nS)\gamma$  where ( $n = 1, 2, 3$ ), a channel with much lower sensitivity than the  $H \rightarrow b\bar{b}$  to the Yukawa coupling to b-quarks.

More recently the ATLAS collaboration has searched for another quarkonia final state where the Higgs boson decays to  $\phi\gamma$  [180] at the LHC Run 2 and a center-of-mass energy of 13 TeV, with a specific trigger. This channel could probe deviations from the strange-quark Yukawa coupling of the Higgs boson. Its sensitivity is several orders of magnitude above the expectation from the SM Higgs boson. Other quarkonia final states, such as the  $\rho\gamma$ , which could potentially probe the Yukawa coupling to light quarks, can also be searched for.

### III.8.6. Rare decays outlook

Rare decays such as those described in the above sections have a clearly limited sensitivity. They however already deliver interesting messages. For example, if the coupling of the Higgs boson was as strong in the dimuon channel as it is for the top quark, this mode would have been observed already with large significance. The observed Higgs boson couplings are manifestly non-universal. Further developing these rare decay modes is an important component of the high luminosity program of the LHC to directly probe the couplings of the Higgs boson, and to potentially measure the Yukawa coupling of the Higgs boson to fermions of the second generation, in particular to muons.

### III.9. Searches for non-standard model decay channels

The main decay and production properties of the observed Higgs boson are consistent with predictions of the SM. It may however have other decay channels beyond those anticipated in the SM. Among these and of great interest are the invisible decays into stable particles that interact very weakly with the detector, and that are undetected, such as Dark Matter particle candidates. Other non standard decay

**Table 11.8:** Summary of the channels searched for and the corresponding 95% CL limits from ATLAS and CMS on the branching fraction for the Higgs boson decay to invisible particles assuming a SM Higgs boson production cross section. The results in parentheses are the expected exclusions. [\*] indicates analyses based only on 8 TeV data. When combining Run 1 results and the results from the  $\approx 2 \text{ fb}^{-1}$  of 13 TeV data acquired in 2015, the CMS observed (expected) limit improves to  $< 32$  (26) % at 95% CL.

	ATLAS (Run 1)	CMS (Run 1)	CMS (13 TeV, 2015)
ggF (monojet); $H \rightarrow \text{inv.}$	–	67 (71) % [*]	–
VBF; $H \rightarrow \text{inv.}$	28 (31) %	57 (40) % [*]	69 (62) %
$Z \rightarrow \ell^+ \ell^-$ ; $H \rightarrow \text{inv.}$	75 (62) %	75 (91) %	125 (125) %
$Z \rightarrow b\bar{b}$ ; $H \rightarrow \text{inv.}$	–	182 (189) % [*]	–
$Z \rightarrow jj$ ; $H \rightarrow \text{inv.}$	78 (86) %	–	–
Combination of all direct searches	25 (27) %	36 (30) %	–

channels that have been investigated are the decays of the Higgs particle to hidden valley or dark particles.

### III.9.1. Invisible decays of the Higgs boson

The discovery of the Higgs boson immediately raised the question of its couplings to dark matter and how it could be used to reveal its existence at colliders, using the Higgs boson as a portal to dark matter (see Ref. [181] and references therein). If kinematically accessible and with a sufficiently large coupling to the Higgs boson, dark matter particles, such as, e.g., neutralinos in SUSY models, graviscalars in models with extra dimensions or heavy neutrinos in the context of four-generation fermion models, would manifest themselves as invisible decays of the Higgs boson, thus strongly motivating searches for the invisible decays of the Higgs boson.

To identify an invisibly decaying Higgs boson at the LHC, it must be produced in association with other particles. Searches for invisible decays of the Higgs particle at the LHC have been carried out in three associated production modes of the Higgs boson with the highest SM cross sections and target events with large missing energy.

The ggF production mode has the largest SM cross section but it usually results in the Higgs boson being created alone and hence leaving no characteristic signature in the detector of its invisible decay. One way to search for invisible decays in ggF production mode is to look for events with the “monojet” topology arising from initial state gluon radiation and containing missing energy. The major irreducible background in such searches stems from  $Z + \text{jets}$  events where the  $Z$  boson decays into a pair of neutrinos. The analysis with the best sensitivity targets the VBF production topology but suffers from large backgrounds arising from events with two jets and large missing energy. The VH mode has much smaller cross section but the presence of a W or Z boson allows a variety of final states that can be tagged with relatively low background.

ATLAS [182–185], and CMS [186–189] have searched for such final states but have observed no significant excess over predicted backgrounds. Table 11.8 summarizes the 95% CL limits on the invisible decays of the Higgs boson assuming SM Higgs boson production cross section and corresponding detector acceptances.

### III.9.2. Exotic Higgs boson decays

The 125 GeV Higgs boson not only serves as a probe for potential dark matter candidates, but also to search for other exotic particles arising from fields associated with a low-mass hidden sector. Such hidden sectors are composed of fields that are singlets under the SM group  $\text{SU}(3) \times \text{SU}(2) \times U(1)$ . These models are referred to as hidden valley models [190, 191]. Since a light Higgs boson is a particle with a narrow width, even modest couplings to new states can give rise to a significant modification of Higgs phenomenology through exotic decays. Simple hidden valley models exist in which the Higgs boson decays to an invisible fundamental particle, which has a long lifetime to decay back to SM particles through small mixings with the SM

Higgs boson; Ref. [191] describes an example. The Higgs boson may also decay to a pair of hidden valley “v-quarks,” which subsequently hadronize in the hidden sector, forming “v-mesons.” These mesons often prefer to decay to the heaviest state kinematically available, so that a possible signature is  $H \rightarrow 4b$ . Some of the v-mesons may be stable, implying a mixed missing energy plus heavy flavor final state. In other cases, the v-mesons may decay to leptons, implying the presence of low mass lepton resonances in high- $H_T$  events [192]. Other scenarios have been studied [193] in which Higgs bosons decay predominantly into light hidden sector particles, either directly, or through light SUSY states, and with subsequent cascades that increase the multiplicity of hidden sector particles. In such scenarios, the high multiplicity hidden sector particles, after decaying back into the SM, appear in the detector as clusters of collimated leptons known as lepton jets.

A variety of models have been investigated searching for final states involving dark photons and hidden valley scalars. The resulting topologies searched for are prompt electron jets in the WH production process [194], displaced muonic jets [195], four muons final state, and long lived weakly interacting particles [196]. The latter occur not only in hidden valley scenarios, but also in gauge-mediated extensions of the minimal supersymmetric standard model (MSSM), the MSSM with R-parity violation, and inelastic dark matter [197]. Finally the CMS collaboration has performed a search for pair production of light bosons [198]. Such a scenario can occur in supersymmetric models with additional hidden (or dark) valleys.

## IV. Combining the main channels

As described in Section II, there are five main production modes of a SM Higgs boson at the LHC. In the LHC Run 1 dataset the predicted numbers of SM Higgs bosons produced per experiment are approximately 0.5 million, 40,000, 20,000 and 3,000 in the gluon fusion, vector boson fusion, the associated VH and  $t\bar{t}H$  production modes respectively<sup>3</sup>. There are also five main decay channels: the  $\gamma\gamma$ ,  $ZZ$ ,  $WW$ ,  $\tau^+\tau^-$  and  $b\bar{b}$ . Analyses using exclusive categories according to production modes have been designed to maximize the sensitivity of the analyses to the presence of a signal using known characteristic features of these modes. These categories can also be used to further separate production modes for each decay channel. The typical number of events selected eventually in each decay channel ranges from a fraction of an event to  $\mathcal{O}(100)$  events per experiment.

The analysis strategy used by the LHC and Tevatron experiments to perform the searches for the Higgs boson has been based on the Higgs decay modes. It is a natural choice given that it focusses on the decay products of the object searched for. However, for each channel, exclusive subchannels have been defined according to the Higgs production processes and in the results presented these subchannels have been combined. The natural extension of this approach in order to probe further the production and decay modes of the Higgs boson is to combine the analysis channels together. Such a combination is also used in Section VI to further measure the coupling properties of the Higgs boson.

At the LHC or the Tevatron, the total cross section cannot be measured in any of the production modes. As a consequence, neither the absolute branching fractions nor the total natural width of the Higgs boson can be directly measured. However, a combined measurement of the large variety of categories described in Section III, with different sensitivities to various production and decay modes permits a wide variety of measurements of the production, decay or in general coupling properties. These measurements require, in general, a limited but nevertheless restrictive number of assumptions.

In this section, results will be given combining not only different channels, but also the ATLAS and CMS results together [141]. These

<sup>3</sup> Similarly at the Tevatron where the CDF and D0 experiments have gathered approximately  $10 \text{ fb}^{-1}$  of data at 1.96 TeV, the predicted numbers of SM Higgs boson events produced per experiment are approximately 10,000 and 2,000 events in the gluon fusion and VH associated production, respectively.

**Table 11.9:** Summary of the main production categories used in the analysis channels involved in the combined measurement of the coupling properties of the Higgs boson. (A) and (C) indicate respectively when ATLAS and CMS include the specific exclusive category in the combination. For the ttH channel, note that the number of leptons ( $\ell$ ) indicates typically the type of decay of the two top quarks.

	$\gamma\gamma$	ZZ ( $4\ell$ )	WW ( $\ell\nu\ell\nu$ )	$\tau^+\tau^-$	$b\bar{b}$
ggF (high $p_T^H$ )	A	A	—	A	—
ggF (incl. or low $p_T^H$ )	A - C	A - C	A - C	—	—
ggF 1-jet	—	C	A - C	C	—
VBF	A - C	A - C	A - C	A - C	C
WH (1- $\ell$ )	A - C	A	A - C	C	A - C
WH (two jets)	A - C	A - C	A - C	—	—
ZH (0- $\ell$ )	A - C	A	—	—	A - C
ZH (2- $\ell$ )	A - C	A	A - C	C	A - C
ZH (two jets)	A - C	A - C	A - C	—	—
ttH (1- $\ell$ )	A - C	—	A - C	A - C	A - C
ttH (2- $\ell$ )	—	—	A - C	A - C	A - C
ttH (hadronic)	A - C	—	—	—	A

results were derived by the two collaborations, taking rigorously into account all correlations in the systematic uncertainties and in the large number of channels and their categories. This combination has led the two collaborations to a more precise experimental portrait of the Higgs boson. This work also concludes and synthesizes the analyses of the main production and decay channels of the Higgs boson at the Run 1 of the LHC.

In this section, only the results on the main Higgs boson production and decay modes will be discussed. The combination framework described herein will also be used in Section VI, to discuss the measurements of the coupling properties of the Higgs boson.

#### IV.1. Principles of the combination

The combination of the Higgs boson analysis channels in each experiment and for the two experiments together is done using a fit of a signal and background model to the data. As described above the data is made of a large number of categories, aiming at reconstructing exclusive production and decay modes. In the combination of ATLAS and CMS [141] there are approximately 600 categories. The combination is a simultaneous fit to all these categories, using a reduced number of parameters of interest and the Higgs boson mass fixed at its measured value (see Section III.2). A synoptic view of the main production categories is illustrated in Table 11.9. The much larger number of categories present in the ATLAS and CMS combination [141], is due to additional separation in terms of finer exclusive production regions, decay channels of the Z and the W bosons, and taus, control regions where little-to-no signal is present, and different center-of-mass energies. It should be noted that the individual combination performed by ATLAS [199] included two additional decay channels: the  $\mu^+\mu^-$  and  $Z\gamma$ , for the sake of simplicity these channels were omitted in the ATLAS-CMS combination. In addition, a  $H \rightarrow b\bar{b}$  analysis performed by CMS [13] and included in its own combination, has been omitted from the ATLAS-CMS combination.

The key to understanding how the combination of channels works relies on the combination master formula, which expresses for each category, indexed by  $c$ , of a given channel (typically a category covers mostly one decay mode, but possibly various production modes), the measured number of signal events  $n_s^c$  as a function of a limited number

of parameters as follows:

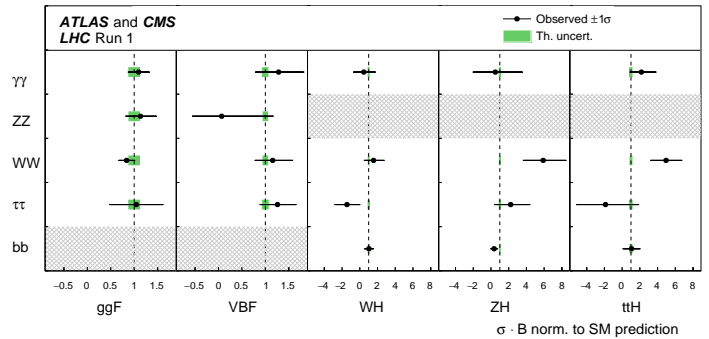
$$n_s^c = \left( \sum_{i,f} \mu_i \sigma_i^{SM} \times A_{if}^c \times \varepsilon_{if}^c \times \mu_f \text{BR}_f^{SM} \right) \times \mathcal{L}^c \quad (11.13)$$

The production index is defined as  $i \in \{ggH, VBF, VH, ttH\}$  and the decay index is defined as  $f \in \{\gamma\gamma, WW, ZZ, b\bar{b}, \tau\tau\}$  while  $\sigma_i^{SM}$  and  $\text{BR}_f^{SM}$  are the corresponding production cross sections and decay branching fractions, estimated as described in Section II, assuming that the Higgs boson is that of the SM.  $A_{if}^c$  and  $\varepsilon_{if}^c$  are the signal acceptance and the reconstruction efficiency for given production and decay modes in the category  $c$ .  $\mathcal{L}^c$  is the integrated luminosity used for that specific category. For the purpose of this review, these parameters can be considered as fixed<sup>4</sup>.

The parameters of interest in the master formula are the signal strength parameters  $\mu_i$  and  $\mu_f$ . It is important to note that the formula relies on the factorization of the production cross section and decay branching fraction, which assumes the narrow width approximation. The width of the Higgs boson will be discussed in Section V, however for the precision needed here, the fact that the Higgs boson has been observed in decay channels with high mass resolution as a resonance is sufficient to validate this hypothesis. It is also manifest in the above equation that the ten parameters for the production modes ( $\mu_i$ ) and decay modes ( $\mu_f$ ) cannot be determined simultaneously. This illustrates that total cross sections or branching fractions cannot be measured without further assumptions in this fit.

The master formula also illustrates an important caveat to the measurement of signal strength parameters. In case these are interpreted as scale factors of the production cross sections or branching fractions, then all the other quantities such as the acceptances and efficiencies,  $A_{if}^c$  and  $\varepsilon_{if}^c$ , need to be assumed as independent and fixed to their estimated values for the SM Higgs boson. An additional important caveat to note concerning these combined results is that only the normalizations are varied, while the discriminating variables for the signal are not modified and are still used in the fit. These caveats are of particular importance in the use of the combination to measure the coupling properties of the Higgs boson as discussed in Section VI. For relatively small perturbations of the couplings of the Higgs boson from the SM values, this hypothesis is valid.

However the 25 products,  $\mu_i \times \mu_f$ , can be considered as free parameters and in principle measurable (if there is sufficient sensitivity from specific categories). Measuring the products of signal strengths can be viewed as the measurements of the cross sections times the branching fraction,  $\sigma \cdot \text{BR}$ . The results are reported in Table 11.10 for the combination of ATLAS and CMS and they are illustrated in Fig. 11.10.



**Figure 11.10:** Combined measurements of the products  $\sigma \cdot \text{BR}$  for the five main production and five main decay modes.

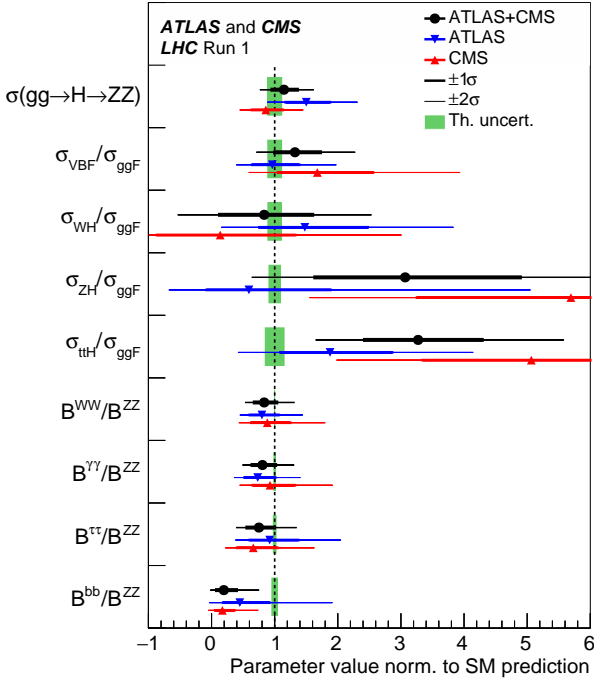
<sup>4</sup> In the combination performed by the ATLAS and CMS experiments the systematic uncertainties on these parameters are taken into account by allowing these parameters to vary in the fit.

It is remarkable that of the 25 possible combinations of production and decay modes in the main channels, the fit to ATLAS and CMS data allows the measurement of 20. A coherent picture emerges with an excellent consistency between the observation in each channel and the expectation for a SM Higgs boson.

This 20 parameter fit quantifies, with very little theoretical input, the current experimental knowledge of the main production and decays modes. It is also a very useful tool to further understand the influential channels in the measurements of the Higgs couplings. Without a loss of independence of the theoretical predictions, a less general fit allowing for further interpretations concerning production cross sections and branching fractions is possible. In this fit a reference process, measured with high precision and significance, is used to parametrize all the other processes. In the ATLAS-CMS combination [141] the gluon fusion production mechanism in the  $H \rightarrow ZZ$  decay mode is chosen. Then, the master formula applies with the following parameters for all  $i$  and  $f$  indices except when both  $i = ggF$  and  $f = ZZ$

$$\begin{aligned} \mu_i &= \frac{\sigma_i}{\sigma_{ggF} \sigma_i^{SM}} \times \mu_{gg \rightarrow H \rightarrow ZZ} \\ \mu_f &= \frac{BR_i}{BR_{ZZ} BR_i^{SM}} \\ \mu_{gg \rightarrow H \rightarrow ZZ} &= \mu_{ggF} \times \mu_{ZZ} \times \sigma_i^{SM} \times BR_f^{SM} \end{aligned} \quad (11.14)$$

The result of the combination with these 9 parameters is illustrated in Fig. 11.11. It allows interesting conclusions on the production and decay properties of the Higgs boson. It shows that the ratio of the decay rates to Z and W bosons is as expected from the SM, a direct illustration of the custodial symmetry, and also quantifies the relative precision at which the  $t\bar{t}H$  coupling is currently measured. It also shows that with the improved precision stemming from the combination, no significant deviations from the SM is observed.



**Figure 11.11:** Measurement of the  $\sigma(gg \rightarrow H \rightarrow ZZ)$  cross section and of the ratios of cross sections and branching fractions from the combination of the ATLAS and CMS measurements. The results from each experiment are also shown. The results are normalised to the SM predictions for the various parameters. The shaded bands indicate the theoretical uncertainties in these predictions.

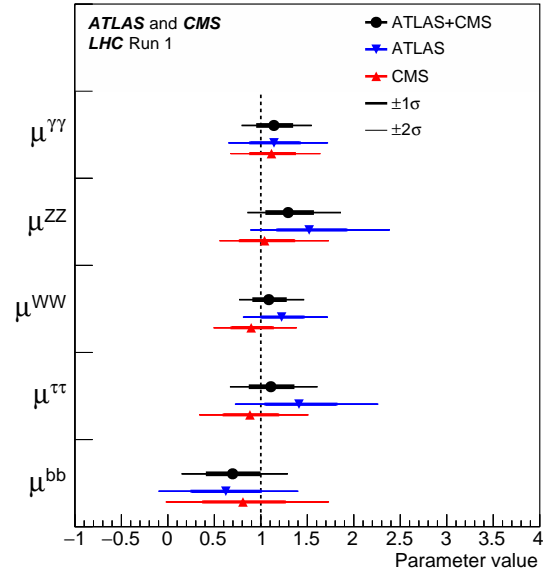
Finally, the most constrained fit in this combination, and historically the first made, allows for only one single parameter to vary *i.e.*  $\forall(i, f), \mu_i = \mu_f = \mu$ . This global signal strength model provides the most precise and simple probe of the compatibility of the signal with the SM Higgs boson. This model is sensitive to any deviation from the SM Higgs boson couplings provided that these deviations do not cancel overall. The combined global signal strength is

$$\mu = 1.09 \pm 0.07 (\text{stat}) \pm 0.04 (\text{expt}) \pm 0.03 (\text{th. bkg}) \pm 0.07 (\text{th. sig})$$

This overall signal strength is fully compatible with the SM expectation of 1, with a precision of 10%. It is interesting to note that the major uncertainties in this first measurement arises from the limited precision in the theoretical predictions for the signal production processes.

#### IV.2. Characterization of the main decay modes and observation of Higgs decays to taus

Despite the large number of decay channels, since the cross sections cannot be independently measured, from the measurements described in this section it is impossible to measure decay branching fractions without a loss of generality. The simplest assumption that can be made is that the production cross sections are those of the SM Higgs boson, which is equivalent to assuming that for all  $i$  indices  $\mu_i = 1$ . All branching fractions  $\mu_f$  can then be measured in a simple 5 parameter fit. The result of this fit is illustrated in Fig. 11.12, and the measured signal strengths are reported in Table 11.11.



**Figure 11.12:** The signal strengths  $\mu$  measured by the ATLAS and CMS experiments [141] for the five principal Higgs boson decay channels assuming that the production cross sections are those of the SM.

Table 11.11 also reports the results of a similar combination by each experiment from their data. For the main discovery modes  $\gamma\gamma$ ,  $ZZ$  and  $WW$ , the combined significance is not computed as these decay modes have been firmly established by each experiment independently. However for the  $\tau^+\tau^-$  and  $b\bar{b}$  decay modes these results shed new combined light on the observation significance in these channels.

For the  $\tau^+\tau^-$  channel, ATLAS and CMS are both sensitive and have observed excesses in their data. The individual results are not sufficiently significant to claim an observation, but combined they are. This conclusion can be made also in a more generic manner using the ratio of branching fractions model described above. It should be noted that in the search for  $H \rightarrow \tau\tau$  decay, the most sensitive production mode is the VBF process, the experimental evidence for which is discussed in Section IV.3.

**Table 11.11:** Summary of the significances of the excesses observed for the main decay processes. The  $\gamma\gamma$ ,  $ZZ$ , and  $W^+W^-$  decay modes have been established at more than  $5\sigma$  by both the ATLAS and CMS experiments individually, the combined observation significance therefore exceeds  $5\sigma$  and is not reported here.

	Expected Z	Observed Z
$\gamma\gamma$	4.6 $\sigma$ (ATLAS) 5.3 $\sigma$ (CMS)	5.2 $\sigma$ (ATLAS) 4.6 $\sigma$ (CMS)
$ZZ$	6.2 $\sigma$ (ATLAS) 6.3 $\sigma$ (CMS)	8.1 $\sigma$ (ATLAS) 6.5 $\sigma$ (CMS)
$WW$	5.9 $\sigma$ (ATLAS) 5.4 $\sigma$ (CMS)	6.5 $\sigma$ (ATLAS) 4.7 $\sigma$ (CMS)
$\tau^+\tau^-$	3.4 $\sigma$ (ATLAS) 3.9 $\sigma$ (CMS)	4.5 $\sigma$ (ATLAS) 3.8 $\sigma$ (CMS)
$b\bar{b}$	2.6 $\sigma$ (ATLAS) 2.5 $\sigma$ (CMS)	1.4 $\sigma$ (ATLAS) 2.1 $\sigma$ (CMS)
$\tau^+\tau^-$ (Comb.)	5.0 $\sigma$	5.5 $\sigma$
$b\bar{b}$ (Combined)	3.7 $\sigma$	2.6 $\sigma$

**Table 11.12:** Summary of the combined significance of observation for the main production processes. The ggF process has been established at more than  $5\sigma$  by both the ATLAS and CMS experiments individually, the combined observation significance far exceeds  $5\sigma$  and is not reported here.

	Expected Z	Observed Z
ggF	Ind. Obs.	Ind. Obs.
VBF	4.6 $\sigma$	5.4 $\sigma$
WH	2.7 $\sigma$	2.4 $\sigma$
ZH	2.9 $\sigma$	2.3 $\sigma$
VH	4.2 $\sigma$	3.5 $\sigma$
ttH	2.0 $\sigma$	4.4 $\sigma$

As illustrated in Table 11.11, ATLAS and CMS are both much less sensitive to the  $H \rightarrow b\bar{b}$  decay mode. The available sensitivity comes mostly from the VH process, as discussed earlier in this section. The combined significance of  $3.7\sigma$  is sufficient to suggest evidence, however ATLAS and CMS observations are both low with respect to the rate expected for the SM Higgs boson. With the increased production cross sections at 13 TeV and the much larger dataset expected, this channel will undoubtedly be followed with great attention at the Run 2 of the LHC.

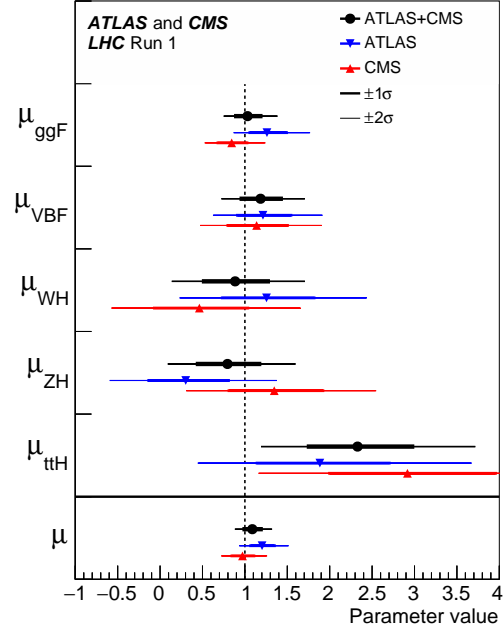
#### IV.3. Characterization of the main production modes and evidence for VBF production

As discussed earlier, most analysis channels are divided into several exclusive categories allowing for an increased overall sensitivity and to measure the various Higgs production modes. The cross sections of the main production modes can be measured assuming that the branching fractions are those of the SM Higgs boson, *i.e.* for all  $f$  indices  $\mu_f = 1$ . These assumptions lead to a 5 parameter combination. The result is illustrated in Fig. 11.13 for the ATLAS-CMS combination [141]. The significance of observation of the production modes are reported in Table 11.12.

The gluon fusion production process is the dominant production mode. Although no numerical estimate of combined significance of observation for this process has been given by the experiments, it is considered as established due to the overwhelming evidence from the three main discovery channels. None of the other production modes have been firmly established by the experiments individually. These show that for the VBF mode, the combination has a large sensitivity and produced a combined observation of  $5.4\sigma$ , establishing this process with a rate compatible with that expected from the SM Higgs boson. A similar conclusion can be reached but with assumptions from the fit to the ratio  $\sigma_{VBF}/\sigma_{ggF}$  discussed earlier in this section.

It is interesting to note that despite the low sensitivity to the  $ttH$  production mode, the excesses observed in several  $ttH$

channels (discussed in Section III.6.1), lead to a significance of direct observation for  $ttH$  production in excess of  $4\sigma$ . The compatibility of this observation with the SM production rate is at the  $2.3\sigma$  level. Given the increased sensitivity expected at the higher center-of-mass energy of 13 TeV, a great attention will undoubtedly be devoted to this channel.



**Figure 11.13:** The signal strengths  $\mu$  measured by the ATLAS and CMS experiments [141] in the five principal production modes and their combination, assuming that the decay branching fractions are those of the SM Higgs boson.

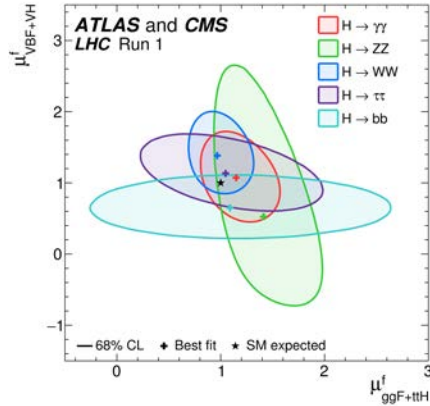
One can further reduce the number of parameters of interest by grouping the ggF and  $ttH$  production modes together, as they both originate in the SM from the Yukawa coupling to the top quark and the VH and VBF processes together as they both originate from the coupling of the Higgs boson to massive vector bosons. Grouping here means that the two production cross sections are scaled by a single parameter of interest, which is equivalent to fixing their ratio to the expected ratio for a SM Higgs boson. The 2-dimensional profile of the likelihood for each main decay channel individually is shown in Fig. 11.14, illustrating the relative constraint that each decay channel is imposing on the production process and the correlation between the grouped processes. This correlation stems from the cross contamination of the exclusive categories with processes from the two groups. Concerning the VBF production mode, these results show its manifest predominance in the  $\tau^+\tau^-$  channel and the importance of the  $\gamma\gamma$  and WW channels in constraining it. While in most cases the  $ggF + ttH$  group is mostly constrained by the indirect gluon fusion process, in the case of the  $b\bar{b}$  channel, the bulk of the constraint comes from the  $ttH$  process.

### V. Main quantum numbers and width of the Higgs boson

#### V.1. Main quantum numbers $J^{PC}$

Probing the Higgs boson quantum numbers is essential to further unveiling its coupling properties. The measurements of the signal event yields of the observed new state in all the channels discussed in Sections III and IV and their compatibility with the SM Higgs boson predictions, give a qualitative, but nonetheless compelling indication of its nature. This qualitative picture is further complemented by the implications of the observation of the particle in the diphoton channel. According to the Landau–Yang theorem [200], the observation made





**Figure 11.14:** Two dimensional likelihood contours for individual production signal strengths for the  $VBF + VH$  versus  $ggF + ttH$  processes for various Higgs boson decay modes for the ATLAS and CMS experiment combination.

in the diphoton channel excludes the spin-1 hypothesis and restricts possibilities for the spin of the observed particle to 0 or 2.

The Landau–Yang theorem does not apply if the observed state is not decaying to a pair of photons but to a pair of scalars subsequently decaying to two very collimated pairs of photons (as for example in the case of  $H \rightarrow a_1 a_1 \rightarrow 4\gamma$ ). This possibility has not been rigorously excluded but is not experimentally favored since tight selection criteria are applied on the electromagnetic shower shapes of the reconstructed photons. A more systematic analysis of shower shapes and the fraction of conversions could be performed to further discriminate between the single prompt photon and the two overlapping photons hypotheses. There are also potential theoretical loopholes concerning the applicability of the Landau–Yang theorem [200], such as off-shell vector boson decays. However, for the observed particle not to be of spin 0 and +1 parity would require an improbable conspiracy of effects. It is nevertheless important that this hypothesis be independently tested, in particular since the measurements of coupling properties of the Higgs boson assume that the observed state is CP-even.

#### V.1.1. Charge conjugation

The charge conjugation quantum number is multiplicative, therefore given that the Higgs-like particle is observed in the  $H \rightarrow \gamma\gamma$  channel, and given that photons are C-odd eigenstates, assuming C conservation, the observed neutral particle should be C-even.

#### V.1.2. Spin and parity

To probe the spin and parity quantum numbers of the discovered particle, a systematic analysis of its production and decay processes is performed in several analyses, designed to be independent of the event yields measured and relying instead on the production and the decay angles, and on the threshold distributions as long as a significant signal is observed (*i.e.* an excess over the expected background that can be used to further discriminate between signal hypotheses) of the produced particle. These analyses are based on probing various alternative models of spin and parity. These models can be expressed in terms of an effective Lagrangian [201] or in terms of helicity amplitudes [202, 203]. The two approaches are equivalent. In the following, the effective Lagrangian formalism is chosen to describe the models considered and a restricted number of models are discussed [201]. In the analysis performed by CMS [202] a larger number of models have been investigated, however the main channels studied by both experiments are essentially the same and the main conclusions are similar and fully consistent.

##### (i) Spin-0 model

The interaction Lagrangian relevant for the analysis of spin-0 particle interaction with a pair of W- or Z-boson with either fixed or mixed SM and BSM CP-even couplings or CP-odd couplings, is the following:

$$\begin{aligned} \mathcal{L}_0^{W,Z} \supset & \left\{ \cos(\alpha) \kappa_{SM} \left[ \frac{1}{2} g_{HZZ} Z_\mu Z^\mu + g_{HWW} W_\mu^+ W^{-\mu} \right] \right. \\ & - \frac{1}{4\Lambda} [\cos(\alpha) \kappa_{HZZ} Z_{\mu\nu} Z^{\mu\nu} + \sin(\alpha) \kappa_{AZZ} Z_{\mu\nu} \tilde{Z}^{\mu\nu}] \\ & \left. - \frac{1}{2\Lambda} [\cos(\alpha) \kappa_{HWW} W_{\mu\nu}^+ W^{-\mu\nu} + \sin(\alpha) \kappa_{AWW} W_{\mu\nu}^+ \tilde{W}^{-\mu\nu}] \right\} H_0 \end{aligned} \quad (11.15)$$

Where  $V^\mu = Z^\mu, W^{+\mu}$  are the vector boson fields,  $V^{\pm\mu\nu}$  are the reduced field tensors and  $\tilde{V}^{\pm\mu\nu} = 1/2 \epsilon^{\mu\nu\rho\sigma} V_{\rho\sigma}$  are the dual tensor fields. Here,  $\Lambda$  defines an effective theory energy scale. The factors  $\kappa_{SM}, \kappa_{HZZ}, \kappa_{HWW}, \kappa_{AZZ}, \kappa_{AWW}$  denote the coupling constants corresponding of the coupling of the SM, BSM CP-even and CP-odd components of the Higgs field  $H_0$  to the W and Z fields. The mixing angle  $\alpha$  allows for the production of CP-mixed state and the CP-symmetry is broken when  $\alpha \neq 0, \pi$ .

This formalism can be used to probe both CP-mixing for a spin-0 state or specific alternative hypotheses such as a pure CP-odd state ( $J^P = 0^-$ ) corresponding to  $\alpha = \pi/2$ ,  $\kappa_{SM} = \kappa_{HVV} = 0$  and  $\kappa_{AVV} = 1$ . A BSM CP-even state  $J^P = 0^+_h$  corresponds to  $\alpha = 0$ ,  $\kappa_{SM} = \kappa_{AVV} = 0$  and  $\kappa_{HVV} = 1$ . These hypotheses are compared to the SM Higgs boson hypothesis corresponding to  $\alpha = 0$  and  $\kappa_{HVV} = \kappa_{AVV} = 0$  and  $\kappa_{SM} = 1$ .

##### (ii) Spin-2 model

The graviton inspired interaction Lagrangian for a spin-2 boson  $X^{\mu\nu}$  for a color, weak and electromagnetic singlet spin-2 resonance uniquely interacting with the energy momentum tensor  $\mathcal{T}^{V,f}$  of vector bosons  $V$  or fermions  $f$ , can be written as follows [204]:

$$\mathcal{L}_2 \supset \frac{1}{\Lambda} \left[ \sum_V \xi_V \mathcal{T}_{\mu\nu}^V X^{\mu\nu} + \sum_f \xi_f \mathcal{T}_{\mu\nu}^f X^{\mu\nu} \right]$$

Where the strength of the interaction is determined by the couplings  $\xi_V$  and  $\xi_f$ . The simplest scenario, referred to as the universal couplings (UC), corresponds to  $\xi_V = \xi_f$ . These models predict a large branching ratio to photons (of approximately 5%) and negligible couplings to massive gauge bosons ( $W$  and  $Z$ ). Such scenarios are therefore disfavored and other models are investigated where the couplings of the  $W$ ,  $Z$  and  $\gamma$  are assumed to be independent. Universality of the couplings refers to  $\xi_g = \xi_q$ . Two other scenarios are considered with low light-quark fraction where  $\xi_q = 0$  and the low gluon-fraction where  $\xi_q = 2\xi_g$ . In these scenarios a large enhancement of the tail of the transverse momentum of the spin-2 state is expected and requires a further selection requirement in order to probe the models within the range of validity of the effective field theory. Two requirements are considered,  $p_T^X < 300$  GeV and  $p_T^X < 125$  GeV [201].

#### V.1.3. Probing fixed $J^P$ scenarios

At the LHC, the determination of the spin and CP properties of the Higgs boson is done independently from the total rates measurement, it uses a global angular helicity analysis and, when applicable, the study of threshold effects. The channels used for this analysis,  $H \rightarrow \gamma\gamma$ ,  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  and  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$ , are those where the observation of a signal is unambiguous.

At the Tevatron, an analysis using the threshold distribution of the production of the discovered state [205] in the associated production mode  $VH$  with subsequent decay to a pair of  $b$ -quarks was performed by the D0 collaboration.

##### (i) The $VH$ production at D0

The mass of the  $VH$  system is a powerful discriminant to distinguish a  $J^P = 0^+_h$ , with a threshold behavior in  $d\sigma/dM^2 \sim \beta$  from  $0^-$  or  $2^+$  with threshold behaviors respectively in  $\sim \beta^3$  and  $\sim \beta^5$  (for a graviton like spin 2) [205]. The  $VH$  mass observable, not only discriminates signal hypotheses, but also has an increased separation between the  $0^-$  and  $2^+$  hypotheses with respect to the backgrounds, thus allowing, with a small and not yet significant signal yield, to exclude that the observed state is  $0^-$  at 98% CL [206] and  $2^+$  at the 99.9% CL [207].

### (ii) The $\gamma\gamma$ channel at the LHC

In the  $H \rightarrow \gamma\gamma$  channel, the analysis is performed inclusively using the production angle  $\cos\theta_{CS}^*$  and the transverse momentum of the diphoton pair [201]. The definition chosen for the polar angle in the rest frame is the Collins–Soper frame, which is defined as the bisector axis of the momenta of the incoming protons in the diphoton rest frame. The SM Higgs signal distribution is expected to be uniform with a cutoff due to the selection requirements on the photons transverse momentum. The  $H \rightarrow \gamma\gamma$  channel is mostly sensitive to the gluon-initiated spin-2 production scenarios, which yield a  $\cos\theta_{CS}^*$  distribution peaking at values close to 1. The limits are derived from a fit of the signal in bins of  $\cos\theta_{CS}^*$  and diphoton transverse momentum and are summarized in Fig. 11.15 for ATLAS, only combined results are shown. The data shows a good compatibility with the SM  $0^+$  hypothesis and contributes strongly to the exclusion of several Spin-2 scenarios. The conclusions are the same from CMS results [202].

### (iii) The $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$ channel at the LHC

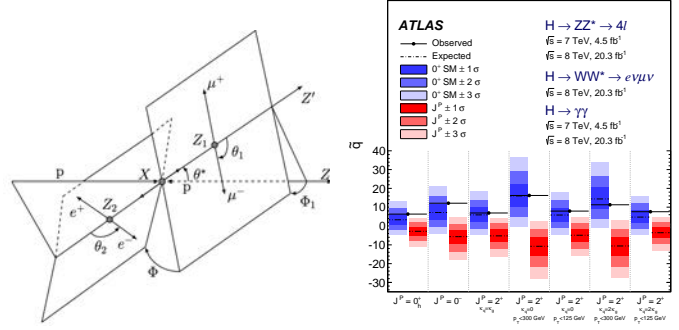
In the  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  channel, the production and decay angles cannot be easily reconstructed due to the presence of neutrinos in the final state, however sensitivity arises from the V-A structure of the decay of the  $W$  bosons. A scalar state thus yields a clear spin correlation pattern that implies that the charged leptons  $e$  or  $\mu$  from the decays of the  $W$  bosons are produced close to one another in the transverse plane. This feature, which impacts observables such as the azimuthal angle between the two leptons  $\Delta\Phi_{\ell\ell}$  or their invariant mass  $M_{\ell\ell}$  in addition of the threshold behavior of the decay which is used in kinematic variables such as the transverse mass defined in Section III, can be used to discriminate between various spin and parity hypotheses. The approach adopted by ATLAS uses a multivariate discriminant, whereas CMS uses a 2D-fit of the dilepton mass and the transverse mass. The results of the  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  analyses alone are summarized in Fig. 11.15 for ATLAS and in combination with other channels. Spin-1 hypotheses ( $1^+$  and  $1^-$ ) have also been tested with this channel by ATLAS and CMS. ATLAS and CMS exclude the  $1^+$  and  $1^-$  hypotheses at more than 95% CL.

### (iv) The $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$ channel at the LHC

The  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  coupling analysis, as described in Section III, also uses a discriminant based on the  $0^+$  nature of the Higgs boson to further discriminate the signal from the background. In this analysis this feature is used to discriminate between signal hypotheses. The observables sensitive to the spin and parity are [208] the masses of the two  $Z$  bosons (due to the threshold dependence of the mass of the off-shell  $Z$  boson), two production angle  $\theta^*$  and  $\phi_1$ , and three decay angles,  $\phi$ ,  $\theta_1$  and  $\theta_2$ . The production and decay angles defined as:

- $\theta_1$  and  $\theta_2$ , the angles between the negative final state lepton and the direction of flight of  $Z_1$  and  $Z_2$  in the rest frame.
- $\phi$ , the angle between the decay planes of the four final state leptons expressed in the four lepton rest frame.
- $\phi_1$ , the angle defined between the decay plane of the leading lepton pair and a plane defined by the vector of the  $Z_1$  in the four lepton rest frame and the positive direction of the proton axis.
- $\theta^*$ , the production angle of the  $Z_1$  defined in the four lepton rest frame with respect to the proton axis.

These angles are illustrated in Fig. 11.15. There are two approaches to this analysis. The first, used by CMS, is a matrix element likelihood approach where a kinematic discriminant is defined based on the ratio of the signal and background probabilities. These probabilities are defined using the leading-order matrix elements. A similar approach is also performed by ATLAS as a cross check of their main result. The main approach adopted by ATLAS is the combination of sensitive observables with a Boosted Decision Tree. These analyses are sensitive to various  $J^P$  hypotheses and in particular discriminate the  $0^+$  hypothesis from the  $0^-$ . In all scenarios investigated and for both the ATLAS and CMS experiments, the data are compatible with the  $0^+$  hypothesis. ATLAS [203] and CMS [202] exclude a pseudoscalar nature of the observed boson at CL<sub>S</sub> levels of 98% and 99.8%.

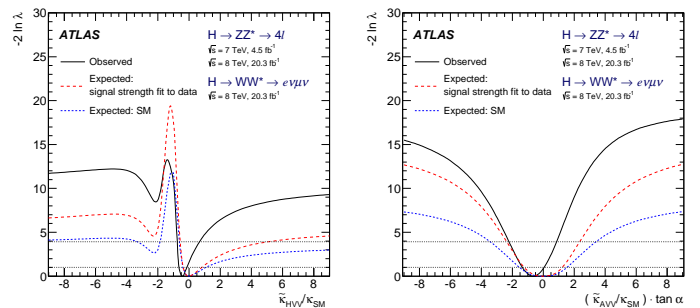


**Figure 11.15:** (Left) Definition of the production and decay angles defined for the  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  final state [202]. Expected distributions of the test statistic for the SM hypothesis (in blue) and several alternative spin and parity hypotheses (in red).

#### V.1.4. Probing anomalous HVV couplings

The careful study of the kinematic properties of the events observed in the  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  and  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  channel, and in particular the angular distributions described above, allows one to further probe the HVV coupling beyond testing fixed hypotheses. Assuming that the observed particle is a spin-0 state, and using several discriminating observables in the  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  and  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  channels, the anomalous terms in the formalism of Eq. (11.15) can be probed. In the approach of helicity amplitudes used by CMS [202], all terms are essentially equivalent, except for one additional phase which is neglected in Eq. (11.15).

Results are derived in terms of the parameters  $\tilde{\kappa}_{HVV} = v/\Lambda \kappa_{HVV}$  and  $\tilde{\kappa}_{AVV} = v/\Lambda \kappa_{AVV}$ , and more precisely as measurements of  $\tilde{\kappa}_{HVV}/\kappa_{SM}$  and  $\tan\alpha \cdot \tilde{\kappa}_{AVV}/\kappa_{SM}$  as shown in Fig. 11.16. These parameters can be interpreted as mixing parameters of a tensor anomalous CP-even coupling and a CP-odd component. The measurements are made in the  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  and  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  channels independently and then combined assuming that the  $\tilde{\kappa}_{HVV}/\kappa_{SM}$  and  $\tan\alpha \cdot \tilde{\kappa}_{AVV}/\kappa_{SM}$  are the same for the  $W$  and  $Z$  vector bosons. Only the combination of the  $WW$  and  $ZZ$  channels is shown in Fig. 11.16. The asymmetric shape of the likelihood as a function of  $\tilde{\kappa}_{HVV}/\kappa_{SM}$  is mainly due to the interference between the BSM and the SM contributions that give a maximal deviation from the SM predictions for negative relative values of the BSM couplings. In Fig. 11.16 the expected likelihood profiles for a SM Higgs boson are also displayed. While no significant deviation from the SM Higgs boson expectation is observed, the precision of the measurements of the mixing parameters is fairly low. The results from the CMS measurements [202] are very similar and the conclusions the same.



**Figure 11.16:** Likelihood profiles for the  $\tilde{\kappa}_{HVV}$  and  $\tilde{\kappa}_{AVV} \cdot \tan\alpha$  parameters, representing respectively CP-even and CP-odd anomalous couplings of the Higgs boson.

### V.2. Off-shell couplings of the Higgs boson

In the dominant gluon fusion production mode, the production cross section of a off-shell Higgs boson is known to be sizable. This follows as a consequence of the enhanced couplings of the Higgs boson to the longitudinal polarizations of the massive vector bosons at high energy.

The off-shell to on-shell cross section ratio is approximately 8% in the SM. Still the Higgs contribution to  $VV$  production at large invariant mass remains small compared to the background. It is nevertheless interesting to probe Higgs production in this regime as it is sensitive to new physics beyond the SM.

The difficulty in the off-shell  $VV$  analysis, beyond the small signal-to-background ratio, is due to a large negative interference between the signal and the  $gg \rightarrow VV$  background. boson signal in the far off-shell domain results in a deficit of events with respect to the expectation from background only events. It is only when the off-shell couplings of the Higgs boson are larger than expected in the SM that the presence of a signal appears as an excess over the background expectation. One additional intricacy arises from the precision in the prediction of the rate for  $gg \rightarrow VV$ , a loop process at lowest order, and its interference with the signal. At the time of the publications of the results from the ATLAS [209] and CMS [210] a full NLO prediction had not been computed.

It is interesting to note that in this regime the Higgs boson is studied as a propagator and not as a particle. The measurement of its off-shell couplings is therefore absolute and does not rely on the knowledge of the total Higgs boson width. The off-shell couplings constraints can then be used to indirectly constrain the natural width of the Higgs boson, under specific assumptions detailed in Section V.3.3.

This measurement has been carried out in the  $H \rightarrow ZZ \rightarrow 4\ell$ ,  $H \rightarrow ZZ \rightarrow \ell\ell\nu\nu$  and  $H \rightarrow WW \rightarrow \ell\nu\ell\nu$  channels. To enhance the sensitivity of the analysis the knowledge of the full kinematics of the events is important. In particular the signal and the background can be further distinguished by the invariant mass of the  $VV$  system, which is more accurately accessible in the  $H \rightarrow ZZ \rightarrow 4\ell$  channel. Angular distributions also play an important role in this analysis. For these reasons the  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  channel is significantly more sensitive than  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$ . The CMS results in Refs. [210] and [211] include the VBF and VH processes through the selection of two additional jets in the final state. The ATLAS results do not have a specific selection for the VBF or VH production processes, but their contributions are taken into account.

Limits on the off-shell rates have been reported for the two channels by ATLAS [209] and CMS [211]. The combined results assuming that the off-shell rates in the  $ZZ$  and  $WW$  channels scale equally, are given for two different hypotheses on the VBF production rate: fixing it to its SM value or scaling it as the gluon fusion rate. The observed (expected) limits on the off-shell rate fraction with respect to its SM expectation is 6.7 (9.1) for ATLAS [209] with the VBF rate fixed to its SM value and 2.4 (6.2) for CMS [211] where no assumption is made on the relative production rates of gluon-fusion and VBF. In both cases the custodial symmetry is assumed and the ratio of the rates in the  $ZZ$  and  $WW$  decays are fixed to those of the Standard Model. Results without this assumption have also been reported in Ref.[211].

### V.3. The Higgs boson width

In the SM, the Higgs boson width is very precisely predicted once the Higgs boson mass is known. For a mass of 125.1 GeV, the Higgs boson has a very narrow width of 4.2 MeV. It is dominated by the fermionic decays partial width at approximately 75%, while the vector boson modes are suppressed and contribute 25% only.

At the LHC or the Tevatron, in all production modes, only the cross sections times branching fractions can be measured. As a consequence, the total natural width of the Higgs boson cannot be inferred from measurements of Higgs boson rates. Direct constraints on the Higgs boson width are much larger than the expected natural width of the SM Higgs boson.

**Table 11.13:** The observed (expected) direct 95%CL constraints on the natural width of the 125 GeV resonance from fits to the  $\gamma\gamma$  and  $ZZ$  mass spectra and to the  $4\ell$  vertex lifetime.

	$M_{\gamma\gamma}$ mass spectrum	$M_{4\ell}$ spectrum	$4\ell$ vertex lifetime
ATLAS	$< 5.0(6.2)$ GeV	$< 2.6(6.2)$ GeV	—
CMS	$< 2.4(3.1)$ GeV	$< 3.4(2.8)$ GeV	$> 3.5 \times 10^{-12}$ GeV

#### V.3.1. Direct constraints

Analysis of the reconstructed mass lineshape in the two channels with a good mass resolution, the  $H \rightarrow \gamma\gamma$  and  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$ , allow for a direct measurement of the width of the SM Higgs boson. The intrinsic mass resolution in these channels is about 1-2 GeV, much larger than the expected width of the SM Higgs boson. As a result only upper limits on the Higgs boson width have been measured by ATLAS [132] and CMS [133]. The two main challenges of direct constraints on the width through the measurement of the lineshape are: (i) the modeling of resolution uncertainties and (ii) the modeling of the interference between the signal and the continuum background which can be sizable for large widths, in particular in the range where direct constraints are set. Given that these interference effects are small with respect to the individual channels sensitivity, they are neglected in deriving constraints on the total width. The combined constraints however, being more precise could be affected by the interference. ATLAS [132] has therefore not combined the constraints on the width from the two channels. The results are reported in Table 11.13. These constraints are still three orders of magnitude larger than the expected SM width and are fully compatible with the SM hypothesis.

Another direct constraint on the Higgs boson width can be obtained in the  $H \rightarrow Z^{(*)}Z^{(*)} \rightarrow 4\ell$  channel, from the measurement of the average lifetime of the Higgs boson calculated from the displacement of the four-lepton vertex from the beam spot. This analysis has been carried out by CMS [210], using the measured decay length. The measured  $c\tau_H$  is  $2_{-2}^{+25} \mu\text{m}$ , yielding an observed (and expected) limit at the 95%CL of  $c\tau_H < 57(56) \mu\text{m}$ . From this upper limit on the lifetime of the Higgs boson. The 95% CL lower limit on its natural width is  $\Gamma_H > 3.5 \times 10^{-12}$  GeV.

#### V.3.2. Indirect constraints from mass shift in the diphoton channel

In the diphoton channel, it was noticed in [212], that the effect of the interference between the main signal  $gg \rightarrow H \rightarrow \gamma\gamma$  and the continuum irreducible background  $gg \rightarrow \gamma\gamma$ , taking into account detector resolution effects, is responsible for a non negligible mass shift. The size of the mass shift depends on the total width of the Higgs boson and it was suggested that measuring this mass shift could provide a constraint on the width [212]. Comparing the mass measured in the diphoton channel with the mass measured in the four-leptons channel is subject to non negligible detector calibration systematic uncertainties, however it was further noticed that the mass shift has a dependence also on the diphoton transverse momentum. The total width of the Higgs boson could therefore be constrained using the diphoton channel alone.

Further studies were performed by the ATLAS collaboration to estimate the size of the expected mass shift [213]. The expected shift in mass in the diphoton channel is  $35 \pm 9$  MeV for the SM Higgs boson. Very preliminary studies of the sensitivity of this method to estimate the width of the Higgs boson in the high-luminosity regime have been made by ATLAS [214] and yield an expected 95% CL upper limit on the total width of approximately 200 MeV from  $3 \text{ ab}^{-1}$  of 14 TeV data.

#### V.3.3. Indirect constraints from off-shell couplings

Using simultaneously on-shell and off-shell measurements in the  $VV$  channels, it was noticed [215] that the total width of the Higgs could be constrained. This can be illustrated from the parametrization of the signal strength measurements both on-shell ( $\mu_{\text{on-shell}}$ ) and off-shell ( $\mu_{\text{off-shell}}$ ) as a function of the couplings modifiers  $\kappa_g$  and  $\kappa_V$  parameterizing the main process  $gg \rightarrow H \rightarrow VV$ . The on-shell signal



strength can be written as:

$$\mu_{\text{on-shell}} = \frac{\kappa_{g,\text{on-shell}}^2 \kappa_{V,\text{on-shell}}^2}{\Gamma_H/\Gamma_{SM}}$$

while in the case of the off-shell signal strength where the Higgs boson is a propagator:

$$\mu_{\text{on-shell}} = \kappa_{g,\text{off-shell}}^2 \kappa_{V,\text{off-shell}}^2$$

Then, with the following assumption

$$\kappa_{g,\text{on-shell}}^2 \kappa_{V,\text{on-shell}}^2 \leq \kappa_{g,\text{off-shell}}^2 \kappa_{V,\text{off-shell}}^2,$$

from the on-shell and off-shell constraints. This assumes that no new physics alters the Higgs boson couplings in the off-shell regime, i.e. that the running of its couplings is negligible in the off-shell regime. Both ATLAS [209] and CMS [210,211] have used their off-shell production limits to constrain the width of the Higgs boson.

Both ATLAS and CMS analyses use the kinematic event characteristics to further gain in sensitivity to discriminate between the signal and background. The ATLAS analysis assumes that there are no anomalous couplings of the Higgs boson to vector bosons, and obtains 95% CL observed (expected) upper limit on the total width of  $5.7 \times \Gamma_{SM}$  ( $9.0 \times \Gamma_{SM}$ ) [209]. In the CMS approach, results are also derived allowing for anomalous couplings of the Higgs boson, therefore reducing the discriminating power of the kinematic variables used in the analysis but reducing the model dependence. When standard couplings are assumed for the Higgs boson, the observed (expected) limit on the total width is  $6.2 \times \Gamma_{SM}$  ( $9.8 \times \Gamma_{SM}$ ) for the  $ZZ$  channel only [210]. Without assumptions on the anomalous couplings of the Higgs boson, the observed (expected) limit on the total width is  $10.9 \times \Gamma_{SM}$  ( $17.4 \times \Gamma_{SM}$ ) [210].

The CMS experiment has also combined the  $ZZ$  and  $W^+W^-$  channels while keeping the gluon-fusion and VBF production processes separate. For the gluon fusion mode the observed (expected) combined upper limit at the 95% CL on the total width of the Higgs boson is  $2.4 \times \Gamma_{SM}$  ( $6.2 \times \Gamma_{SM}$ ) [211], while for the VBF production mode the exclusion limits are  $19.3 \times \Gamma_{SM}$  ( $34.4 \times \Gamma_{SM}$ ) [211].

ATLAS has also performed a study of the prospects for measuring the Higgs width in the four lepton channel alone, in the high luminosity phase of the LHC for and projects that for a LHC luminosity of  $3 \text{ ab}^{-1}$ , the width of the Higgs boson could be measured with the following precision [216]:

$$\Gamma_H = 4.2_{-2.1}^{+1.5} \text{ MeV}$$

## VI. Probing the coupling properties of the Higgs boson

As discussed in Section II, within the SM, all the Higgs couplings are fixed unambiguously once all the particle masses are known. Any deviation in the measurement of the couplings of the Higgs boson could therefore signal physics beyond the SM.

Measuring the Higgs couplings without relying on the SM assumption requires a general framework treating deviations from the SM coherently at the quantum level in order to provide theoretical predictions for relevant observables to be confronted with experimental data. The first attempt in that direction was the development of the so-called  $\kappa$ -formalism where the SM Higgs couplings are rescaled by factors  $\kappa_f$ , keeping the same Lorentz structure of the interactions. This formalism allows for simple interpretation of the signal strengths  $\mu$  measured in the various Higgs channels and it has been used to test various physics scenarios, like the existence of additional new particles contributing to the radiative Higgs production and decays, or to probe various symmetries of the SM itself, in particular the custodial symmetry. But the  $\kappa$ -formalism has obvious limitations and certainly does not capture the most general deformations of the SM, even under the assumptions of heavy and decoupling new physics. A particularly acute shortcoming at the time Higgs physics is entering a precision era

is the lack of proficiency of the  $\kappa$ 's to assert the richness of kinematical distributions beyond simple signal strength measurements. Several extensions and alternative approaches are being developed as part of the activities of the Higgs cross-section working group.

The Higgs Pseudo-Observable (HPO) approach [217] is providing a particularly elegant formalism to report the data in terms of a finite set of on-shell form factors parametrizing amplitudes of physical processes subject to constraints from Lorentz invariance and other general requirements like analyticity, unitarity, and crossing symmetry. These form factors are expanded in powers of kinematical invariants of the process around the known poles of SM particles, assuming that poles from BSM particles are absent in the relevant energy regime. A set of HPOs have been proposed to characterize both the Higgs decays and the EW Higgs production channels, thus exploring different kinematical regimes. Prospective studies concluded that these HPOs can be measured/bounded at the percent level at the HL-LHC and could therefore be used to constrain some explicit models of New Physics.

Another promising approach to characterize the possible Higgs coupling deviations induced by physics beyond the SM is the use of Effective Field Theories (EFT). This approach assumes again that the new physics degrees of freedom are sufficiently heavy to be integrated out and they simply give rise to effective interactions among the light SM particles. By construction the effective Lagrangians cannot account for deviations in Higgs physics induced by light degrees of freedom, unless they are added themselves as extra fields in the effective Lagrangians. In Section VII, several examples of models with light degrees of freedom affecting Higgs production and decay rates will be presented. The main advantage of EFTs is their prowess to relate different observables in different sectors and at different energies to constrain a finite set of effective interactions among the SM degrees of freedom. In an EFT, the SM Lagrangian is extended by a set of higher-dimensional operators, and it reproduces the low-energy limit of a more fundamental UV description. It will be assumed that the Higgs boson is part of a CP-even EW doublet. This is motivated by the apparent relation between the Higgs couplings and the masses of the various particles which naturally follows under this assumption of a linear realization of the  $SU(2)_L \times U(1)_Y$  symmetry of the SM. There have been some recent attempts to write the most general EFT bypassing this assumption, see for instance [218].

### VI.1. Effective Lagrangian framework

The EFT has the same field content and the same linearly-realized  $SU(3)_C \times SU(2)_L \times U(1)_Y$  local symmetry as the SM. The difference is the presence of operators with canonical dimension  $D$  larger than 4. These are organized in a systematic expansion in  $D$ , where each consecutive term is suppressed by a larger power of a high mass scale. Assuming baryon and lepton number conservation, the most general Lagrangian takes the form

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \sum_i c_i^{(6)} \mathcal{O}_i^{(6)} + \sum_j c_j^{(8)} \mathcal{O}_j^{(8)} + \dots \quad (11.16)$$

The list of dimension-6 operators was first classified in a systematic way in Ref. [219] after the works of Ref. [220]. Subsequent analyses pointed out the presence of redundant operators, and a minimal and complete list of operators was finally provided in Ref. [221]<sup>5</sup>. For a single family of fermions, there are 76 real ways to deform the SM generated by 59 independent operators (with the 3 families of fermions of the SM, flavor indices can be added to these 59 operators, and furthermore, new operator structures, that have been dismissed by means of Fierz transformations in the single family case, have to be considered, for a total of 2499 real deformations [223]). Of particular interest are the 17 CP-invariant operators, in addition to 8 dipole operators, that affect, at tree-level, the Higgs production and decay rates [224–226]. A convenient list of these operators can be found in

<sup>5</sup> Complete classifications of  $D=8$  operators have recently appeared in the literature, see Ref. [222]. Still, in this review, the EFT Lagrangians will be truncated at the level of dimension-6 operators.

**Table 11.14:** List of 17 CP-even operators affecting, at tree-level, only Higgs production and decay rates (left) as well as EW observables (right). See text for notations.

Operators affecting Higgs physics only	Operators affecting Higgs and EW physics
$\mathcal{O}_r =  \Phi ^2  D_\mu \Phi ^2$	$\mathcal{O}_W = \frac{ig}{2} (\Phi^\dagger \sigma^i \overleftrightarrow{D}^\mu \Phi) (D^\nu W_{\mu\nu})^i$
$\mathcal{O}_6 = \lambda  \Phi ^6$	$\mathcal{O}_B = \frac{ig'}{2} (\Phi^\dagger \overleftrightarrow{D}^\mu \Phi) (\partial^\nu B_{\mu\nu})$
$\mathcal{O}_{BB} = \frac{g'^2}{4}  \Phi ^2 B_{\mu\nu} B^{\mu\nu}$	$\mathcal{O}_T = \frac{1}{2} (\Phi^\dagger \overleftrightarrow{D}^\mu \Phi)^2$
$\mathcal{O}_{WW} = \frac{g^2}{4}  \Phi ^2 W_{\mu\nu}^i W^{i\mu\nu}$	$\mathcal{O}_{HB} = ig' (D^\mu \Phi)^\dagger (D^\nu \Phi) B_{\mu\nu}$
$\mathcal{O}_{GG} = \frac{g_S^2}{4}  \Phi ^2 G_{\mu\nu}^A G^{A\mu\nu}$	$\mathcal{O}_{Hu} = i (\bar{u}_R \gamma^\mu u_R) (\Phi^\dagger \overleftrightarrow{D}_\mu \Phi)$
$\mathcal{O}_{yu} = y_u  \Phi ^2 \bar{q}_L \tilde{\Phi} u_R$	$\mathcal{O}_{Hd} = i (\bar{d}_R \gamma^\mu d_R) (\Phi^\dagger \overleftrightarrow{D}_\mu \Phi)$
$\mathcal{O}_{yd} = y_d  \Phi ^2 \bar{q}_L \Phi d_R$	$\mathcal{O}_{He} = i (\bar{l}_R \gamma^\mu l_R) (\Phi^\dagger \overleftrightarrow{D}_\mu \Phi)$
$\mathcal{O}_{ye} = y_e  \Phi ^2 \bar{L}_L \Phi e_R$	$\mathcal{O}_{Hq} = i (\bar{q}_L \gamma^\mu q_L) (\Phi^\dagger \overleftrightarrow{D}_\mu \Phi)$
	$\mathcal{O}_{Hq}^{(3)} = i (\bar{q}_L \gamma^\mu \sigma^i q_L) (\Phi^\dagger \sigma^i \overleftrightarrow{D}_\mu \Phi)$

**Table 11.15:** List of 8 dipoles operators. See text for notations.

Dipoles operators
$\mathcal{O}_{uB} = g' (\bar{q}_L \tilde{\Phi} \sigma^{\mu\nu} u_R) B_{\mu\nu}$
$\mathcal{O}_{uW} = g (\bar{q}_L \sigma^i \tilde{\Phi} \sigma^{\mu\nu} u_R) W_{\mu\nu}^i$
$\mathcal{O}_{uG} = g_S (\bar{q}_L \tilde{\Phi} \sigma^{\mu\nu} t^A u_R) G_{\mu\nu}^A$
$\mathcal{O}_{dB} = g' (\bar{q}_L \Phi \sigma^{\mu\nu} d_R) B_{\mu\nu}$
$\mathcal{O}_{dW} = g (\bar{q}_L \sigma^i \Phi \sigma^{\mu\nu} d_R) W_{\mu\nu}^i$
$\mathcal{O}_{dG} = g_S (\bar{q}_L \Phi \sigma^{\mu\nu} t^A d_R) G_{\mu\nu}^A$
$\mathcal{O}_{lB} = g' (\bar{L}_L \Phi \sigma^{\mu\nu} l_R) B_{\mu\nu}$
$\mathcal{O}_{lW} = g (\bar{L}_L \sigma^i \Phi \sigma^{\mu\nu} l_R) W_{\mu\nu}^i$

Table 11.14, and Table 11.15. The other operators completing the basis of dimension-6 operators can be found in Ref. [226].

The SM gauge couplings are denoted by  $g', g, g_S$  while  $y_{u,d,e}$  are the SM Yukawa couplings (in the mass eigenstate basis that diagonalizes the general Yukawa coupling matrices  $Y_{u,d,l}$ ) and  $\lambda$  is the SM Higgs quartic coupling. We denote by  $i\Phi^\dagger \overleftrightarrow{D}^\mu \Phi$  the Hermitian derivative  $i\Phi^\dagger (D^\mu \Phi) - i(D^\mu \Phi)^\dagger \Phi$ ,  $\sigma^{\mu\nu} \equiv i[\gamma^\mu, \gamma^\nu]/2$  and  $\tilde{\Phi}$  is the Higgs charge-conjugate doublet:  $\tilde{\Phi} = i\sigma^2 \Phi^*$ . Each operator  $\mathcal{O}_{yu,yd,ye}$  is further assumed to be flavor-aligned with the corresponding fermion mass term, as required in order to avoid large Flavor-Changing Neutral Currents (FCNC) mediated by the tree-level exchange of the Higgs boson. This implies one coefficient for the up-type quarks ( $c_{yu}$ ), one for down-type quarks ( $c_{yd}$ ), and one for the charged leptons ( $c_{ye}$ ), i.e. the  $c_{yu,ud,ye}$  matrices should be proportional to the identity matrix in flavor space.

The choice of the basis of operators is not unique and using the equations of motion, i.e., performing field redefinitions, different dimension-6 operators can be obtained as linear combinations of the operators in the previous tables and of four-fermion operators. Some relations between common bases of operators can be found for instance in Refs. [225, 223]. Different bases have different advantages. For instance the so-called SILH basis [224] better captures the low-energy effects of universal theories in which new physics couples to SM bosons only. The Warsaw basis [221] on the other hand mostly includes vertex corrections and easily connects operators to observables [226]. The basis defined in Table 11.14, and Table 11.15 is particularly well suited for an analysis of the Higgs data. The reason is that the eight operators of the left-hand side of Table 11.14, in the vacuum with  $|\Phi|^2 = v^2/2$ , merely redefine the SM input parameters and therefore were left unconstrained at tree-level before Higgs data are considered. These eight operators modify the physical Higgs vertices and can be probed via the decay processes  $H \rightarrow \gamma\gamma, Z\gamma, b\bar{b}, \tau\tau$  and the production channels  $gg \rightarrow H, VV \rightarrow H, pp \rightarrow t\bar{t}H$  and  $gg \rightarrow HH$ . Section VI.2 illustrates how the Higgs data accumulated at the LHC can (partially) constrain these eight operators, following the initial phenomenological study of Ref. [226]. The other nine operators of Table 11.14 are

tightly constrained by the LEP EW precision measurements (the measurements of the Z-boson couplings to quarks and leptons on the Z-pole) and by diboson production.<sup>6</sup>

The minimal flavor violation assumption imposes Yukawa dependences in the eight dipole operators. For the light generations of fermions, this dependence lowers the induced deviations in the Higgs rates below the experimental sensitivity reachable in any foreseeable future. The corresponding operators in the top sector are not suppressed but they are already constrained by the limit of the top dipole operators imposed by the bounds on the neutron electric dipole moment, on the  $b \rightarrow s\gamma$  and  $b \rightarrow s\ell^+\ell^-$  rates and on the  $t\bar{t}$  cross section [229, 225].

Automatic tools [225, 204] are being developed to analyze the experimental data within an EFT framework.

## VI.2. Probing coupling properties

As described in Section III a framework was developed by the ATLAS and CMS collaboration [141], individually and together, to combine the very large number of exclusive categories aimed at reconstructing the five main decay modes and the five main production modes of the Higgs boson. The general conclusions of this combination in terms of production cross sections and decay modes, illustrating the compatibility of the observation with the expectation from the SM Higgs boson is given in Section III. The same framework with its master formula Eq. (11.13) can be used to further measure coupling properties of the Higgs boson under specific additional assumptions.

### VI.2.1. Combined measurements of the coupling properties of H

#### (i) From effective Lagrangians to Higgs observables

All 8 operators of the effective Lagrangian that were unconstrained before the Higgs data induce, at tree-level, deviations in the Higgs couplings that either respect the Lorentz structure of the SM interactions, or generate simple new interactions of the Higgs boson to the W and Z field strengths, or induce some contact interactions of the Higgs boson to photons (and to a photon and a Z boson) and gluons that take the form of the ones that are generated by integrating out the top quark. In other words, the Higgs couplings are described, in the unitary gauge, by the following effective Lagrangian [230, 44]

$$\begin{aligned}
\mathcal{L} = & \kappa_3 \frac{m_H^2}{2v} H^3 + \kappa_Z \frac{m_Z^2}{v} Z_\mu Z^\mu H + \kappa_W \frac{2m_W^2}{v} W_\mu^+ W^{-\mu} H \\
& + \kappa_g \frac{\alpha_s}{12\pi v} G_{\mu\nu}^a G^{a\mu\nu} H + \kappa_\gamma \frac{\alpha}{2\pi v} A_{\mu\nu} A^{\mu\nu} H + \kappa_{Z\gamma} \frac{\alpha}{\pi v} A_{\mu\nu} Z^{\mu\nu} H \\
& + \kappa_{VV} \frac{\alpha}{2\pi v} \left( \cos^2 \theta_W Z_{\mu\nu} Z^{\mu\nu} + 2 W_{\mu\nu}^+ W^{-\mu\nu} \right) H \\
& - \left( \kappa_t \sum_{f=u,c,t} \frac{m_f}{v} f\bar{f} + \kappa_b \sum_{f=d,s,b} \frac{m_f}{v} f\bar{f} + \kappa_\tau \sum_{f=e,\mu,\tau} \frac{m_f}{v} f\bar{f} \right) H.
\end{aligned} \tag{11.17}$$

The correspondence between the effective coefficients of the dimension-6 operators and the  $\kappa$ 's can be found for instance in Ref. [45]. In the SM, the Higgs boson does not couple to massless gauge bosons at tree level, hence  $\kappa_g = \kappa_\gamma = \kappa_{Z\gamma} = 0$ . Nonetheless, the contact operators are generated radiatively by SM particles loops. In particular, the top quark gives a contribution to the 3 coefficients  $\kappa_g, \kappa_\gamma, \kappa_{Z\gamma}$  that does not decouple in the infinite top mass limit. For instance, in that limit  $\kappa_\gamma = \kappa_g = 1$  [24, 25, 231].

The coefficient for the contact interactions of the Higgs boson to the W and Z field strengths is not independent but obeys the relation

$$(1 - \cos^4 \theta_W) \kappa_{VV} = \sin 2\theta_W \kappa_{Z\gamma} + \sin^2 \theta_W \kappa_{\gamma\gamma}. \tag{11.18}$$

This relation is a general consequence of the custodial symmetry [225], which also imposes  $\kappa_Z = \kappa_W$  at leading order ( $\kappa_Z/\kappa_W - 1$  is a measure

<sup>6</sup> There remains an accidental flat direction [227] in the fit of anomalous gauge boson couplings using LEP2 data on diboson production alone. This flat direction can be lifted when LHC Higgs data are considered [228].

of custodial symmetry breaking and as such is already constrained by electroweak precision data and the bounds on anomalous gauge couplings). When the Higgs boson is part of an  $SU(2)_L$  doublet, the custodial symmetry could only be broken by the  $\mathcal{O}_T = \frac{1}{2}(\Phi^\dagger \overleftrightarrow{D}^\mu \Phi)^2$  operator at the level of dimension-6 operators and it is accidentally realized among the interactions with four derivatives, like the contact interactions considered.

The coefficient  $\kappa_3$  can be accessed directly only through double Higgs production processes, hence it will remain largely unconstrained at the LHC. The LHC will also have a limited sensitivity on the coefficient  $\kappa_\tau$  since the lepton contribution to the Higgs production cross section remains subdominant and the only way to access the Higgs coupling is via the  $H \rightarrow \tau^+\tau^-$  and possibly  $H \rightarrow \mu^+\mu^-$  channels. Until the associated production of a Higgs with a pair of top quarks is observed, the Higgs coupling to the top quark is only probed indirectly via the one-loop gluon fusion production or the radiative decay into two photons. However, these two processes are only sensitive to the combinations of couplings  $(\kappa_t + \kappa_g)$  and  $(\kappa_t + \kappa_\gamma)$  and not to the individual couplings. Therefore a deviation in the Higgs coupling to the top quark can in principle always be masked by new contact interactions to photons and gluons (and this is precisely what is happening in minimal incarnations of composite Higgs models). The current limited sensitivity in the  $t\bar{t}H$  channel leaves elongated ellipses in the direction  $\kappa_g = \kappa_\gamma = 1 - \kappa_t$ .

The operators already bounded by EW precision data and the limits on anomalous gauge couplings modify in general the Lorentz structure of the Higgs couplings and hence induce some modifications of the kinematical differential distributions [232, 233]. A promising way to have a direct access to the effective coefficients of these operators in Higgs physics is to study the  $VH$  associated production with a  $W$  or a  $Z$  at large invariant mass of the  $VH$  system [232, 234]. It has not been estimated yet whether the sensitivity on the determination of the effective coefficients in these measurements can compete with the one derived for the study of anomalous gauge couplings. In any case, these differential distributions could also be a way to directly test the hypothesis that the Higgs boson belongs to an  $SU(2)_L$  doublet together with the longitudinal components of the massive electroweak gauge bosons.

### (ii) Interpretations of the experimental data

The measurements of the coupling properties of the Higgs boson are entirely based on the formalism of the effective Lagrangian described in Section V.6.2.i. Measurements of coupling properties in this framework means measurements of the parameters of the model Eq. (11.17) or combinations of these parameters with different sets of assumptions.

These measurements are carried out with the combination framework described in Section IV where the  $\mu_i$  and  $\mu_f$  signal strength parameters are further interpreted in terms of modifiers of the SM couplings  $\kappa_k$  where  $k \in \{Z, W, f, g, \gamma, Z\gamma\}$  as in Eq. (11.17). These coupling modifiers  $\kappa$  are fully motivated as leading order coupling scale factors defined such that the cross sections  $\sigma_j$  and the partial decay widths  $\Gamma_j$  associated with the SM particle  $j$  scale with the factor  $\kappa_j^2$  when compared to the corresponding SM prediction. The number of signal events per category for the various production modes are typically estimated at higher orders in the analyses but are scaled by these single LO-inspired factors, thus not taking into account possible intricacies and correlations of these parameters through the higher order corrections. This approximation is valid within the level of precision of current results and their compatibility with the SM expectation.

In this formalism further assumptions are explicitly made: (i) the signals observed in the different search channels originate from a single narrow resonance with a mass of 125 GeV; (ii) similarly to the combination described in Section IV the narrow width approximation is assumed (to allow the decomposition of signal yields); (iii) the tensor structure of the couplings is assumed to be the same as that of a SM Higgs boson. This means in particular that the observed state is assumed to be a CP-even scalar as in the SM.

Loop-level couplings such as the  $gg \rightarrow H$ ,  $H \rightarrow \gamma\gamma$  and  $H \rightarrow Z\gamma$

can either be treated effectively, with The  $\kappa_g$ ,  $\kappa_\gamma$  and  $\kappa_{Z\gamma}$  as free parameters in the fit or these parameters can be expressed in terms of the know SM field content and as a function of the SM coupling modifiers, in the following way:

$$\begin{aligned}\kappa_g^2(\kappa_t, \kappa_b) &= 1.06 \kappa_t^2 - 0.07 \kappa_t \kappa_b + 0.01 \kappa_b^2 \\ \kappa_\gamma^2(\kappa_F, \kappa_V) &= 1.59 \kappa_V^2 - 0.66 \kappa_V \kappa_F + 0.07 \kappa_F^2 \\ \kappa_{Z\gamma}^2(\kappa_F, \kappa_V) &= 1.12 \kappa_V^2 - 0.15 \kappa_V \kappa_F + 0.03 \kappa_F^2\end{aligned}\quad (11.19)$$

The  $\kappa_{Z\gamma}$  parametrization is used only in the ATLAS combined measurements of the coupling properties of the Higgs boson [199]. Neither the  $Z\gamma$  nor the  $\mu^+\mu^-$  channels are included in the CMS [133] and the ATLAS-CMS combinations [141], which therefore do not use the  $\kappa_{Z\gamma}$  or  $\kappa_\mu$  parameters explicitly. The parametrizations are given for a Higgs boson mass hypothesis of 125.09 GeV (and in the last two expressions, all the Higgs-fermion couplings are assumed to be rescaled by an universal multiplicative factor  $\kappa_F$ ). It can be noted from the expression of  $\kappa_\gamma$  that the coupling of the Higgs boson to photons is dominated by the loop of  $W$  bosons, and it is affected by the top quark loop mostly through its interference with the  $W$  loop. The sensitivity of the current measurements to the relative sign of the fermion and vector boson couplings to the Higgs boson is due to this large negative interference term. The  $\kappa_g$  parameter is expressed in terms of the scaling of production cross sections and therefore also depends on the  $pp$  collisions centre-of-mass energy. The parametrizations of  $\kappa_\gamma$  and  $\kappa_{Z\gamma}$  are obtained from the scaling of partial widths and are only dependent on the Higgs boson mass hypothesis. Experiments use a more complete parametrization with the contributions from the  $b$ -quarks,  $\tau$ -leptons in the loop [230, 44].

The global fit is then performed expressing the  $\mu_i$  and  $\mu_f$  parameters in terms of a limited number of  $\kappa_k$  parameters or their ratios, under various assumptions. The parametrization for the main production modes are:  $\mu_{ggF} = \kappa_g^2$  for the gluon fusion and an effective coupling of the Higgs boson to the gluons;  $\mu_{VBF, VH} = \kappa_V^2$  for the VBF and VH processes when the  $W$  and  $Z$  couplings are assumed to scale equally, and the following expression for the VBF production mode is used:

$$\mu_{VBF}^2(\kappa_W, \kappa_Z) = \frac{\kappa_W^2 \sigma_{WWH} + \kappa_Z^2 \sigma_{ZZH}}{\sigma_{WWH} + \sigma_{ZZH}} \quad (11.20)$$

when the couplings to the  $W$  and  $Z$  bosons are varied independently ( $\sigma_{WWH}$  and  $\sigma_{ZZH}$  denote the VBF cross sections via the fusion of a  $W$  and a  $Z$  boson respectively, the small interference term is neglected);  $\mu_{t\bar{t}H} = \kappa_t^2$  for the  $t\bar{t}H$  production mode. Numerically the production modes signal strengths as a function of the coupling modifiers to the SM fields are:

$$\begin{aligned}\mu_{ggF} &= 1.06 \kappa_t^2 + 0.01 \kappa_b^2 - 0.07 \kappa_t \kappa_b \\ \mu_{VBF} &= 0.74 \kappa_W^2 + 0.26 \kappa_Z^2\end{aligned}$$

The decay mode signal strengths are parametrized as  $\mu_k = \kappa_k^2 / \kappa_H^2$  where  $k \in \{Z, W, f, g, \gamma, Z\gamma\}$  denotes the decay mode and  $\kappa_H$  the overall modifier of the total width. Similarly to the combinations reported in Section IV, when parametrizing signal yields per categories using the combination master formula, it is again manifest that parametrizations as a function of coupling modifiers ( $\kappa$ ) cannot be obtained if  $\kappa_H$  is considered effective, since it is a common factor to all signal yields. However,  $\kappa_H$  can also be treated as an effective parameter or expressed in terms of the coupling modifiers to the SM field content. Its general expression as a function of the Standard Model field content is:

$$\begin{aligned}\kappa_H^2 &= 0.57 \kappa_b^2 + 0.06 \kappa_\tau^2 + 0.03 \kappa_c^2 \\ &+ 0.22 \kappa_W^2 + 0.03 \kappa_Z^2 + 0.09 \kappa_g^2 + 0.0023 \kappa_\gamma^2\end{aligned}\quad (11.21)$$

The general expression of the total width of the Higgs boson can be written as follows:

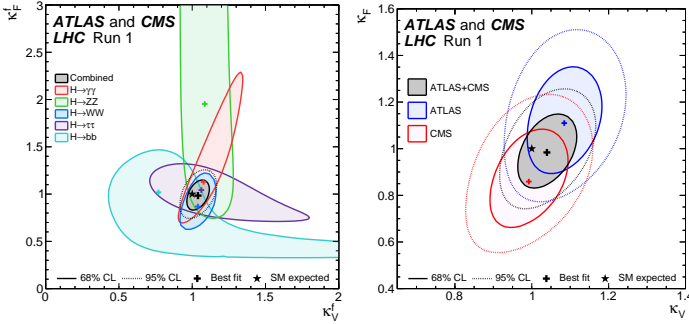
$$\Gamma_H = \frac{\kappa_H^2 \Gamma_H^{SM}}{1 - \text{BR}_{BSM}}$$

where  $\Gamma_H^{SM}$  is the total width of the SM Higgs boson and  $\text{BR}_{\text{BSM}}$  is the branching fraction of the Higgs boson to new particles beyond the SM.

Specific parametrizations will be made in order to address the following aspects of the coupling properties of the Higgs boson under different assumptions: (i) the relative couplings of the Higgs boson to fermions and bosons; (ii) the potential impact of the presence of new particles beyond the SM either in the loops or both in the loops and the decay of the  $H$ ; and (iii) also, more general models either of coupling modifiers or their ratios, under different assumptions.

### (iii) Relative couplings to bosons and fermions

As will be discussed in Section VII.6.3, it is interesting to probe a model where no additional field content is considered in the decay width of the Higgs boson and where the relative couplings of the Higgs boson to  $W$ - and  $Z$ -bosons is fixed to its SM value and where all Yukawa couplings scale with one coupling modifier. In this model only SM particles are assumed to contribute to the gluon fusion and the diphoton loops, all fermion couplings modifiers are required to scale simultaneously with a unique factor  $\kappa_F$  and all vector boson couplings modifiers must scale simultaneously with a unique factor  $\kappa_V$ . This parametrization assumes that no new particles affect the direct decays or the loops. It is a two parameters fit with  $\kappa_V$  and  $\kappa_F$  as parameters of interest. The ATLAS-CMS combined results for each channel independently, the combinations of all channels for the two experiments separately and the results and the overall combination are shown in Fig. 11.17.



**Figure 11.17:** Likelihood contours in the  $(\kappa_F, \kappa_V)$  plane for the ATLAS-CMS combination for the main decay channels separately (left) and for the individual combination of all channels for ATLAS and CMS separately and the complete combined contour (right) [141].

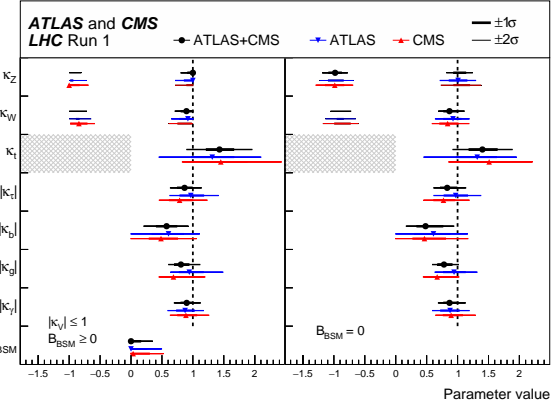
The global fit is only sensitive to the relative sign of  $\kappa_V$  and  $\kappa_F$ . By convention negative values of  $\kappa_F$  can be considered. Such values are not excluded a priori, but would imply the existence of new physics at a light scale and would also raise questions about the stability of such a vacuum [235]. Among the five low mass Higgs channels, only the  $\gamma\gamma$  is sensitive to the sign of  $\kappa_F$  through the interference of the  $W$  and  $t$  loops as shown in Eq. (11.19). The current global fit disfavors a negative value of  $\kappa_F$  at more than five standard deviations. A specific analysis for the Higgs boson production in association with a single top quark has been proposed [236,237] in order to more directly probe the sign of  $\kappa_F$ . All available experimental data show a fair agreement of the SM prediction of the couplings of the Higgs boson to fermions and gauge bosons. The results shown in Fig. 11.17 assume that  $\kappa_F \geq 0$ , however in Ref. [141], a similar combination is done without this assumption. The combined sensitivity to the exclusion of a negative relative sign, is approximately  $5\sigma$  in this model. It is interesting to note that although none of the channels have a significant sensitivity to resolve the sign ambiguity, the combination can, mainly through the  $W-t$  interference in the  $H \rightarrow \gamma\gamma$  channel and the  $H \rightarrow W^+W^-$  channel. The observed exclusion is fully compatible with the expectation [141]. The combined measurements of these

parameters:

$$\begin{aligned}\kappa_V &= 1.04 \pm 0.05 \\ \kappa_F &= 0.98^{+0.11}_{-0.10}\end{aligned}$$

Is already at the 5% level for the  $\kappa_V$  parameter with the Run 1 dataset.

### (iv) Coupling measurements and probing new physics beyond the SM in loops and in the decay



**Figure 11.18:** ATLAS-CMS combined measurements of coupling modifiers.

In the model described above in Section VI.2.1.iii the assumption is that no new fields distort in a perceptible way the loop contributions in the couplings of the  $H$  to gluons and photons and the total width, its couplings to known SM particles are then probed. In a first approach to simultaneously probe new physics beyond the SM in the loops and not in the decay and the couplings of the Higgs boson to SM particles, only one assumption is needed *i.e.* that  $\text{BR}_{\text{BSM}} = 0$ . In this model the coupling of the  $H$  to photons and gluons is effective and  $\kappa_Z$ ,  $\kappa_W$ ,  $\kappa_t$ ,  $|\kappa_\tau|$ , and  $|\kappa_b|$  are measured simultaneously. The absolute value of certain coupling modifiers only indicates the complete degeneracy of combined likelihood for the two signs. It can be noted that when the coupling to gluons is not considered effective, there is some sensitivity to the sign of  $\kappa_b$  through the interference between the top and bottom quarks loops in the gluon fusion process. In this model it is interesting to note that the constraints on the top quark Yukawa coupling comes from the  $ttH$  direct search channels. The expected precision on  $\kappa_t$  is approximately 40%. As discussed in Section III the excesses observed in the  $ttH$  channel yield a large value of  $\kappa_t = 1.40^{+0.24}_{-0.21}$ . The complete set of results from this model is given in Fig. 11.18.

This model, which assumes that no new particles enter the decay of the Higgs boson, also yields very interesting constraints on new physics in the loops through the effective coupling modifiers  $\kappa_g$  and  $\kappa_\gamma$ . The measured values of these parameters:

$$\begin{aligned}\kappa_g &= 0.78^{+0.13}_{-0.10} \\ \kappa_\gamma &= 0.87^{+0.14}_{-0.09}\end{aligned}$$

are fully compatible with the expectation for the SM Higgs boson.

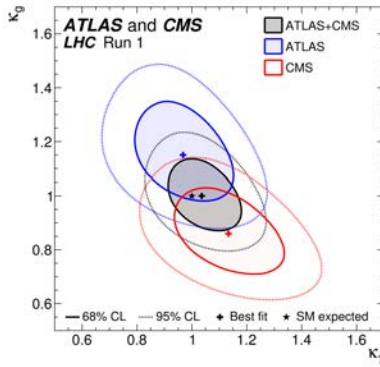
A more constrained model fully focussing on BSM scenarios with new heavy particles contributing to the loops (and not directly in the decays *i.e.*  $\text{BR}_{\text{BSM}} = 0$ ) and where all couplings to the SM particles are assumed to be the same as in the Standard Model ( $\kappa_W = \kappa_Z = \kappa_t = \kappa_b = \kappa_\tau = 1$ ) is also used to constrain the  $\kappa_g$  and  $\kappa_\gamma$  parameters only. The contours of the combined likelihood in the  $(\kappa_\gamma, \kappa_g)$  plane for the ATLAS and CMS experiments and their combination are shown in Fig. 11.19.

This general model requires the strong assumption that the Higgs boson decays only to SM particles. This assumption is necessary due to the degeneracy of solutions given that  $\kappa_H$  is a common factor to all measured signals. The degeneracy can however be resolved

using a constraint on the width of the Higgs boson as the one from the Off-Shell couplings measurements. This approach was used by the ATLAS experiment [199], thus yielding a absolute measurement of the couplings of the Higgs boson.

Another well motivated constraint to resolve the aforementioned degeneracy is unitarity. Simply requiring that  $\kappa_V \leq 1$  allows to free the  $\text{BR}_{\text{BSM}}$  parameter and further probe new physics in the decay of the Higgs boson. An intuitive understanding of how this constraint works can be given by a simple example *e.g.* VBF  $H \rightarrow W^+W^-$  production where the number of signal events will be parametrized by  $(1 - \text{BR}_{\text{BSM}})\kappa_W^4/\kappa_H^2$ , where for a number of signal events observed close to the SM expectation, large values of  $\text{BR}_{\text{BSM}}$  cannot be compensated by a large value of  $\kappa_W$  and is thus limited. Or in other terms, if  $\kappa_W \sim 1$  is preferred from other channels, a low signal in the VBF  $H \rightarrow W^+W^-$  channel would be a sign of the presence of new physics beyond the SM in the Higgs decays. From this general model all the above parameters can be measured in addition to  $\text{BR}_{\text{BSM}}$ . The results of this combination are shown in Fig. 11.18. The results for all parameters do not change significantly with respect to the previous model. A limit can however be set on the beyond the SM branching fraction of the Higgs boson at the 95% CL:

$$\text{BR}_{\text{BSM}} < 34\%$$

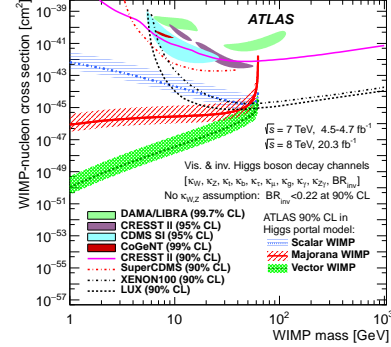


**Figure 11.19:** Likelihood contours of the global fit in the  $(\kappa_g, \kappa_\gamma)$  plane for the ATLAS-CMS combination for the individual combination of all channels for ATLAS and CMS separately and the complete combined contour [141].

In the second approach, new physics is considered also in the decay thus affecting the total width of the  $H$  through decays to particles which are either “invisible” and escape detection in the experiments, or “undetected” which are not distinctive enough to be seen in the current analyses. This approach is complementary to the direct search for invisible decays of the Higgs boson described in Section III. The two approaches can be combined assuming that the undetected branching fraction is negligible. This combination was performed by the ATLAS experiment [238] and yields a limit on the invisible decays of the Higgs boson of  $\text{BR}_{\text{inv}} < 25\%$  at the 95% CL.

This constraint can then be further used to probe Higgs portal models to Dark Matter [239], where an additional weakly interacting particle  $\chi$  with mass typically lower than  $m_H/2$  is introduced as Dark Matter candidate and where the Higgs boson is considered as the only mediator between the SM particles and Dark Matter. In this model it is interesting to express the limit on the invisible branching fraction in terms of strength of interaction of Dark Matter with standard matter, *i.e.* in terms of its interaction cross section with nucleons  $\sigma_{\chi-N}$ . In this model the couplings of the Higgs boson to SM particles are assumed to be those of the SM and the interaction of the Higgs boson with the nucleon is parametrized in Higgs-Nucleon form factor estimated using lattice QCD calculations [239]. The exclusion limits from the constraints on invisible Higgs decays, both direct and indirect from the measurement of the coupling properties of the Higgs boson can be compared to direct detection experiments. For

comparison the limit at 90% CL on the invisible branching fraction of  $\text{BR}_{\text{inv}} < 22\%$  [238] is used and converted into limits on  $\sigma_{\chi-N}$  under several hypotheses on the nature of Dark Matter particles depending mainly on their spin (scalar-, vector- or fermion-like). These results are shown in Fig. 11.20.



**Figure 11.20:** 90% CL upper limits on the WIMP-nucleon scattering cross section as a function of the Dark Matter particle mass. Spin-independent results excluded and favored regions from direct detection experiments are also shown.

### (v) Generic measurement of the $H$ couplings to fermions and gauge bosons

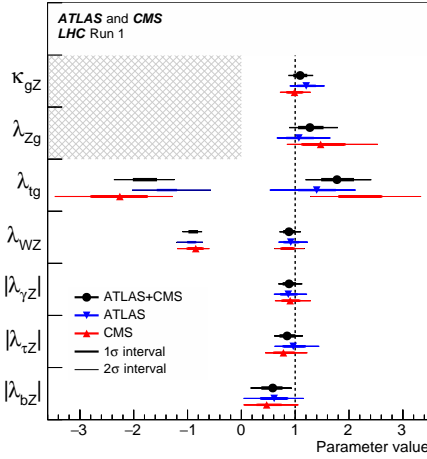
Measuring the couplings of the Higgs boson with requires additional input on constraints on its natural width, or further assumptions. A more generic approach to avoid the degeneracy in the measurement of the coupling modifiers is to probe the coupling properties of the Higgs boson through ratio of couplings. This model, is inspired by the generic model of ratios of cross sections and branching ratios discussed in Section III, where the cross section times branching fraction of the  $gg \rightarrow H \rightarrow ZZ$  process is parametrized as a function of a single coupling modifier:

$$\kappa_{gZ} = \kappa_g \times \frac{\kappa_Z}{\kappa_H}$$

Then all combination signals can be parametrized with the following ratios of coupling modifiers: (i) the  $\lambda_{Zg} = \kappa_Z/\kappa_g$  ratio which is mainly probed by the measurements of the VBF and ZH production; (ii) the  $\lambda_{tg} = \kappa_t/\kappa_g$  ratio constrained by the  $ttH$  production process; (iii) the  $\lambda_{WZ} = \kappa_W/\kappa_Z$  ratio mainly probed by the  $WW$  and  $ZZ$  decay modes; (iv) the  $\lambda_{\tau Z} = \kappa_\tau/\kappa_Z$  ratio constrained by the  $\tau^+\tau^-$  channel; (v) the  $\lambda_{bZ} = \kappa_b/\kappa_Z$  ratio probed mainly by the  $VH(b\bar{b})$  channels; and (vi) the  $\lambda_{\gamma Z} = \kappa_\gamma/\kappa_Z$  ratio constrained by the diphoton channel. In this parametrization the  $ZZ$  channels plays an important normalization role. This model is very general, as it makes neither assumptions on the total width of the Higgs boson or the content of the loops. The results of the combination for this general model are illustrated in Fig. 11.21.

This general model summarizes the status of a flurry of measurements of the Higgs boson that have been carried out at Run 1, in five main production and five main decay modes. Reaching unprecedented level of complexity in a combination with approximately 600 categories and several thousands of parameters for systematic uncertainties. It summarizes the legacy of the measurements of Run 1 and their main conclusions on the coupling properties of the Higgs boson, which can be summarized as follows: (i) the  $\kappa_{gZ}$  parameters shows that the Higgs boson has been firmly observed in a direct  $ZZ$  decay mode; (ii) the  $\lambda_{WZ}$  parameter illustrates the firm observation of the Higgs boson direct decays to the  $W$ -boson and measures directly the ratio of the coupling to the  $W$ - and  $Z$ -bosons a direct probe of the custodial symmetry, already very accurately measured with electroweak precision data; (iii) the  $\lambda_{\gamma Z}$  parameters shows that the coupling of the Higgs boson to photons is compatible with the SM





**Figure 11.21:** ATLAS-CMS combined measurements of ratios of coupling modifiers.

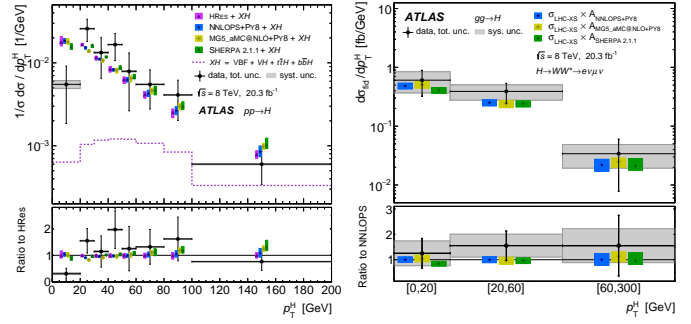
expectation, thus providing a probe of new physics in the decay loop of the Higgs boson to two photons; (iv) the  $\lambda_{\tau Z}$  parameters indicates the evidence in the direct decay of the Higgs boson to a pair of taus of the Yukawa coupling of the Higgs boson to fermions (and in particular to taus); (v) the  $\lambda_{bZ}$  parameter indicates that more data is needed to further establish the Yukawa coupling of the Higgs boson to b-quarks; and (vi) the  $\lambda_{tZ}$  similarly indicates that the more data is needed to further directly constrain the Yukawa coupling of the Higgs boson to top quarks. Each measurement is consistent with the prediction within less than  $2\sigma$ , except for  $\lambda_{bZ}$  and  $\lambda_{tZ}$  which show similar disagreements to those discussed in Section IV. However the probability of the overall compatibility of these measurements with the SM expectation is estimated to be 13%.

### VI.2.2. Differential cross sections

To further characterize the production and decay properties of  $H$ , first measurements of fiducial and differential cross sections have been carried out by the ATLAS collaboration [240], with the 8 TeV dataset of  $pp$  collision at LHC, corresponding to an integrated luminosity of  $20.3\text{ fb}^{-1}$ , in the diphoton channel. The selection criteria to define the fiducial volume are the following: the two highest transverse momentum ( $E_T$ ), isolated final state photons, within  $|\eta| < 2.37$  and with  $105\text{ GeV} < M_{\gamma\gamma} < 160\text{ GeV}$  are selected (the transition region between the barrel and endcap calorimeters is not removed); after the pair is selected, the same cut on  $E_T/M_{\gamma\gamma}$  as in the event selection *i.e.* in excess of 0.35 (0.25) for the two photons is applied. Several observables have been studied: the transverse momentum rapidity of the diphoton system, the production angle in the Collins–Soper frame, the jet multiplicity, the jet veto fractions for a given jet multiplicity, and the transverse momentum distribution of the leading jet. The following additional observables: the difference in azimuthal angle between the leading and the subleading jets, and the transverse component of the vector sum of the momenta of the Higgs boson and dijet system, have also been measured in two jet events. To minimize the model dependence the differential cross sections are given within a specific fiducial region of the two photons. The observables were chosen to probe the production properties and the spin and parity of the  $H$ . The differential cross section in  $H$  transverse momentum is given in Fig. 11.22.

### VI.2.3. Constraints on non-SM Higgs boson interactions in an effective Lagrangian

An example of the possible use of differential cross sections in constraining non-SM Higgs boson couplings in an EFT is given by the ATLAS collaboration [241]. In this analysis, differential cross section measured in the diphoton channel are used to constrain an effective Lagrangian where the SM is supplemented by dimension six CP-even operators of the Strongly Interacting Light Higgs (SILH) formulation and corresponding CP-odd operators. The diphoton differential cross



**Figure 11.22:** Observed differential cross sections in transverse momentum of the  $H$  in the diphoton channel, compared to the prediction of the  $ggF$  process [240].

sections are mainly sensitive to the operators that affect the Higgs boson interactions with gauge bosons and the relevant terms in the effective Lagrangian can be parameterized as:

$$\begin{aligned} \mathcal{L}_{eff} = & \bar{c}_\gamma \mathcal{O}_\gamma + \bar{c}_g \mathcal{O}_g + \bar{c}_{HW} \mathcal{O}_{HW} + \bar{c}_{HB} \mathcal{O}_{HB} + \\ & \tilde{c}_\gamma \tilde{\mathcal{O}}_\gamma + \tilde{c}_g \tilde{\mathcal{O}}_g + \tilde{c}_{HW} \tilde{\mathcal{O}}_{HW} + \tilde{c}_{HB} \tilde{\mathcal{O}}_{HB} \end{aligned} \quad (11.22)$$

Where  $\bar{c}_i$  and  $\tilde{c}_i$  are the effective coefficients corresponding to the CP-even and CP-odd interactions, respectively.

The differential distributions used in this combination are: (i) the transverse momentum of the Higgs boson, (ii) the number of reconstructed jets produced in association with the diphoton pair, (iii) the invariant mass of the diphoton system and (iv) the difference in azimuthal angle of the leading and sub-leading jets in events with two or more jets.

This analysis shows how differential information significantly improves the sensitivity to operators that modify the Higgs boson interaction to photons, gluons and vector bosons both from the main gluon fusion and the vector boson fusion production modes.

## VII. New physics models of EWSB in the light of the Higgs boson discovery

A main theoretical motivation to add a Higgs boson to the SM is that, without it, the longitudinal components of the massive EW gauge bosons would form a strongly coupled system as their scattering amplitude would have grown with their energy, destroying all the predictive power of the model above  $4\pi v \sim 3\text{ TeV}$ . The discovery of a light scalar with couplings to gauge bosons and fermions that are apparently consistent with SM predictions and the slow running of the Higgs self-coupling at high energies allows one to consider the SM as a valid perturbative description of nature all the way to the Planck scale. This picture is admittedly very attractive, but it posits that the Higgs boson is an elementary scalar field, which comes with an intrinsic instability of its mass under radiative corrections. This Higgs naturalness problem calls for new physics around the TeV scale. Supersymmetric models are the most elegant solution to maintain the perturbativity of the SM while alleviating the instability issue. Another possibility is that the Higgs boson itself has a finite size and is composite and thus never feels the UV degrees of freedom that would drag its mass to much higher scales. Both classes of models predict specific modifications from the SM Higgs properties.

The realization of supersymmetry at low energies has many good qualities that render it attractive as a model of new physics. First of all since for every fermion there is a boson of equal mass and effective coupling to the SM-like Higgs, in the case of exact supersymmetry it yields an automatic cancellation of loop corrections to the Higgs mass parameter: (analogous to Eq. (11.2))  $\delta m^2 = 0$  [9, 11]. In practice, it is known that SUSY must be broken in nature since no superpartners of the SM particles have been observed so far. The mass difference between the boson and fermion degrees of freedom is governed by the soft supersymmetry breaking parameters, generically called  $M_{SUSY}$ . Therefore, independently of the precise value of any of the particle

masses and that of its corresponding superpartner, all corrections are proportional to  $M_{SUSY}^2$  times the logarithmic dependence of the ratio of energy scales as in Eq. (11.2). Hence, provided that  $M_{SUSY} \simeq \mathcal{O}(\text{few})$  TeV, the fine-tuning problem is solved, in the sense that the low energy mass parameters of the Higgs sector become insensitive to physics at the GUT or Planck scale. Another interesting feature of SUSY theories is related to the dynamical generation of EWSB [242]. In the SM a negative Higgs mass parameter,  $m^2$ , needs to be inserted by hand to induce EWSB. In SUSY, instead, even if the relevant Higgs mass parameter is positive in the ultraviolet, it may become negative and induce electroweak symmetry breaking radiatively through the strong effect of the top quark-Higgs boson coupling in its renormalization group evolution.

In the following, the Higgs sector will be explored in specific SUSY models. In all of them there is one neutral Higgs boson with properties that resemble those of the SM Higgs boson, whereas additional neutral and charged Higgs bosons are also predicted and are intensively being sought for at the LHC (see Section VII.8). In the simplest SUSY model the lightest Higgs boson mass, that usually plays the role of the SM-like Higgs, is predicted to be less than 135 GeV for stops in the TeV to few TeV range [243] whereas, larger values of the SM-like Higgs boson mass – up to about 250 GeV – can be obtained in non-minimal SUSY extensions of the SM [243]. In general, accommodating a SM-like Higgs boson with mass of 125 GeV results in constraints on the supersymmetric parameter space of specific SUSY models. While naturalness dictates relatively light stops and gluinos, the first and second generation of squarks and sleptons couple weakly to the Higgs sector and may be heavy. Moreover, small values of the  $\mu$  parameter and therefore light Higgsinos would be a signature of a natural realization of electroweak symmetry breaking. Such SUSY spectra, consisting of light stops and light Higgsinos, have been under intense scrutiny by the experimental collaborations [244] in order to derive model-independent bounds on the stop masses and to understand if such natural SUSY scenarios endure [243] and can explain why the Higgs boson remains light.

In the context of weakly coupled models of EWSB one can also consider multiple Higgs  $SU(2)_L$  doublets as well as additional Higgs singlets, triplets or even more complicated multiplet structures, with or without low energy supersymmetry. In general for such models one needs to take into account experimental constraints from precision measurements and flavor changing neutral currents. The LHC signatures of such extended Higgs sectors are largely shaped by the role of the exotic scalar fields in EWSB.

The idea that the Higgs boson itself could be a composite bound state emerging from a new strongly-coupled sector has regained some interest. The composite Higgs idea is an interesting incarnation of EWSB via strong dynamics that smoothly interpolates between the standard Technicolor approach and the true SM limit. To avoid the usual conflict with EW data, it is sufficient if not necessary that a mass gap separates the Higgs resonance from the other resonances of the strong sector. Such a mass gap can naturally follow from dynamics if the strongly-interacting sector exhibits a global symmetry,  $G$ , broken dynamically to a subgroup  $H$  at the scale  $f$ , such that, in addition to the three Nambu–Goldstone bosons of  $SO(4)/SO(3)$  that describe the longitudinal components of the massive  $W$  and  $Z$ , the coset  $G/H$  contains a fourth Nambu–Goldstone boson that can be identified with the physical Higgs boson. Simple examples of such a coset are  $SU(3)/SU(2)$  or  $SO(5)/SO(4)$ , the latter being favored since it is invariant under the custodial symmetry (it is also possible to have non-minimal custodial cosets with extra Goldstone bosons, see for instance Ref. [245]). Attempts to construct composite Higgs models in 4D have been made by Georgi and Kaplan (see for instance Ref. [246]) and modern incarnations have been recently investigated in the framework of 5D warped models where, according to the principles of the AdS/CFT correspondence, the holographic composite Higgs boson then originates from a component of a gauge field along the 5th dimension with appropriate boundary conditions.

A last crucial ingredient in the construction of viable composite Higgs models is the concept of partial compositeness [247], i.e., the idea that there are only linear mass mixings between elementary fields

and composite states<sup>7</sup>. After diagonalization of the mass matrices, the SM particles, fermions and gauge bosons, are admixtures of elementary and composite states and thus they interact with the strong sector, and in particular with the Higgs boson, through their composite component. This setup has important consequences on the flavor properties, chiefly the suppression of large flavor changing neutral currents involving light fermions. It also plays an important role in dynamically generating a potential for the would-be Goldstone bosons. Partial compositeness also links the properties of the Higgs boson to the spectrum of the fermionic resonances, i.e. the partners of the top quark. As in the MSSM, these top partners are really the agents that trigger the EWSB and also generate the mass of the Higgs boson that otherwise would remain an exact Goldstone boson and hence massless. The bounds from the direct searches for the top partners in addition to the usual constraints from EW precision data force the minimal composite Higgs models into some rather unnatural corners of their parameter spaces [15, 249].

### VII.1. Higgs bosons in the minimal supersymmetric standard model (MSSM)

The particle masses and interactions in a supersymmetric theory are uniquely defined as a function of the superpotential and the Kähler potential [250]. A fundamental theory of supersymmetry breaking, however, is unknown at this time. Nevertheless, one can parameterize the low-energy theory in terms of the most general set of soft supersymmetry-breaking operators [243]. The simplest realistic model of low-energy supersymmetry is the minimal supersymmetric extension of the SM (MSSM) [11, 250], that associates a supersymmetric partner to each gauge boson and chiral fermion of the SM, and provides a realistic model of physics at the weak scale. However, even in this minimal model with the most general set of soft supersymmetry-breaking terms more than 100 new parameters are introduced [243]. Fortunately, only a subset of these parameters impact the Higgs phenomenology either directly at tree-level or through quantum effects. Reviews of the properties and phenomenology of the Higgs bosons of the MSSM can be found for example in Refs. [40, 250, 251].

The MSSM contains the particle spectrum of a two-Higgs-doublet model (2HDM) extension of the SM and the corresponding supersymmetric partners. Two Higgs doublets,

$$\Phi_1 = \frac{1}{\sqrt{2}} \begin{pmatrix} \phi_1^0 + i a_1^0 \\ \sqrt{2} \phi_1^- \end{pmatrix}, \quad \Phi_2 = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2} \phi_2^+ \\ \phi_2^0 + i a_2^0 \end{pmatrix}, \quad (11.23)$$

with hypercharge  $Y = -1$  and  $Y = 1$ , respectively, are required to ensure an anomaly-free SUSY extension of the SM and to generate mass for both up-type and down-type quarks and charged leptons [12]. In our notation  $\Phi_{1(2)}$  gives mass to the down(up) type fermions. The Higgs potential reads

$$\begin{aligned} V = & m_1^2 \Phi_1^\dagger \Phi_1 + m_2^2 \Phi_2^\dagger \Phi_2 - m_3^2 (\Phi_1^T i \sigma_2 \Phi_2 + \text{h.c.}) + \frac{1}{2} \lambda_1 (\Phi_1^\dagger \Phi_1)^2 \\ & + \frac{1}{2} \lambda_2 (\Phi_2^\dagger \Phi_2)^2 + \lambda_3 (\Phi_1^\dagger \Phi_1) (\Phi_2^\dagger \Phi_2) + \lambda_4 [\Phi_1^T i \sigma_2 \Phi_2]^2 \\ & + \frac{1}{2} \lambda_5 [(\Phi_1^T i \sigma_2 \Phi_2)^2 + \text{h.c.}] \\ & + [(\lambda_6 (\Phi_1^\dagger \Phi_1) + \lambda_7 (\Phi_2^\dagger \Phi_2)) \Phi_1^T i \sigma_2 \Phi_2 + \text{h.c.}] \end{aligned} \quad (11.24)$$

where  $m_i^2 = \mu^2 + m_{H_i}^2$ , with  $\mu$  being the supersymmetric Higgsino mass parameter and  $m_{H_i}$  (for  $i = 1, 2$ ) the soft supersymmetric breaking mass parameters of the two Higgs doublets;  $m_3^2 \equiv B\mu$  is associated to the B-term soft SUSY breaking parameter; and  $\lambda_i$ , for  $i = 1$  to 7, are all the Higgs quartic couplings. After the spontaneous breaking of the electroweak symmetry, five physical Higgs particles are left in the spectrum: one charged Higgs pair,  $H^\pm$ , one CP-odd neutral

<sup>7</sup> For a pedagogical introduction to models of partial compositeness, see Ref. [248].

scalar,  $A$ , and two CP-even neutral states,  $H$  and  $h$ .

$$\begin{aligned} H^\pm &= \sin\beta\phi_1^\pm + \cos\beta\phi_2^\pm, \\ A &= \sin\beta\operatorname{Im}\phi_1^0 + \cos\beta\operatorname{Im}\phi_2^0, \\ H &= \cos\alpha(\operatorname{Re}\phi_1^0 - v_1) + \sin\alpha(\operatorname{Re}\phi_2^0 - v_2), \\ h &= -\sin\alpha(\operatorname{Re}\phi_1^0 - v_1) + \cos\alpha(\operatorname{Re}\phi_2^0 - v_2), \end{aligned} \quad (11.25)$$

where  $v_i = \langle\phi_i^0\rangle$  for  $i=1,2$  and  $v^2 = v_1^2 + v_2^2 \approx (246\text{ GeV})^2$ . The angle  $\alpha$  diagonalizes the CP-even Higgs squared-mass matrix and is given in terms of the quartic couplings, while  $\beta$  diagonalizes both the CP-odd and charged Higgs sectors with  $\tan\beta = v_2/v_1$ . The  $h$  and  $H$  denote the lightest and heaviest CP-even Higgs bosons, respectively.<sup>8</sup>

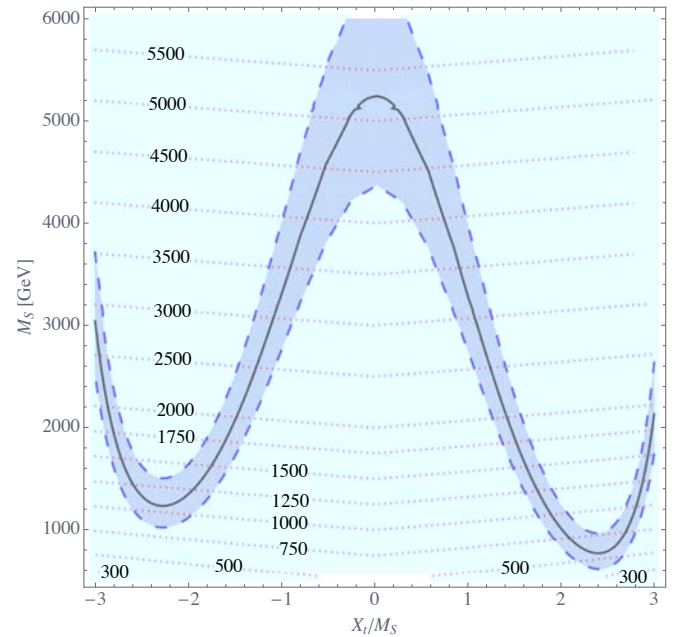
The supersymmetric structure of the theory imposes constraints on the Higgs sector of the model. In particular, at tree level, the parameters of the Higgs self-interaction,  $\lambda_{1,\dots,4}$ , are defined in terms of the electroweak gauge coupling constants, and  $\lambda_{5,6,7} = 0$ . As a result, the Higgs sector at tree level depends on the electroweak gauge coupling constants and the vacuum expectation value  $v$  – or equivalently the  $Z$  gauge boson mass – and is determined by only two free parameters:  $\tan\beta$  and one Higgs boson mass, conventionally chosen to be the CP-odd Higgs boson mass,  $m_A$ . The other tree-level Higgs boson masses are then given in terms of these parameters. In the large  $m_A \gg M_Z$  limit, also called the decoupling limit [252–253],  $\sin\alpha \rightarrow -\cos\beta$ ,  $\cos\alpha \rightarrow \sin\beta$ , hence,  $\cos(\beta - \alpha) \rightarrow 0$  and this implies that the lightest CP-even Higgs  $h$  behaves as the SM Higgs. When  $m_A \geq M_Z$ , the condition  $\cos(\beta - \alpha) \rightarrow 0$  is also called the alignment limit [253–254]. As will be discussed below, in the MSSM the alignment limit can only occur once quantum corrections to the quartic couplings have been included. The tree level value of  $m_h$  is maximized not only for  $m_A \gg M_Z$  but also for  $\tan\beta \gg 1$ . For  $m_A \gg M_Z$  it acquires a maximum value  $m_h = M_Z \cos 2\beta$ .

Radiative corrections have a significant impact on the values of Higgs boson masses and couplings in the MSSM. The dominant radiative effects to the SM-like Higgs mass arise from the incomplete cancellation between top and scalar-top (stop) loops and at large  $\tan\beta$  also from sbottom and stau loops. The loop contributions to the tree level quartic couplings depend on the SUSY spectrum, and render  $\lambda_{5,6,7}$  non zero. The stop, sbottom and stau masses and mixing angles depend on the supersymmetric Higgsino mass parameter  $\mu$  and on the soft-supersymmetry-breaking parameters [11, 250]:  $M_Q$ ,  $M_U$ ,  $M_D$ ,  $M_L$ ,  $M_E$ , and  $A_t$ ,  $A_b$ ,  $A_\tau$ . The first three of these are the left-chiral and the right-chiral top and bottom scalar quark mass parameters. The next two are the left-chiral stau/sneutrino and the right-chiral stau mass parameters, and the last three are the trilinear parameters that enter in the off-diagonal squark/slepton mixing elements:  $X_t \equiv A_t - \mu \cot\beta$  and  $X_{b,\tau} \equiv A_{b,\tau} - \mu \tan\beta$ . At the two-loop level, the masses of the gluino and the electroweak gaugino also enter in the calculations.

Radiative corrections to the Higgs boson masses have been computed using a number of techniques, with a variety of approximations; see references in Ref. [243]. For large  $\tan\beta$ , the stau/sbottom mixing parameters and masses are also relevant. In the large  $m_A$  (decoupling) limit and for  $\tan\beta \gg 1$ , the  $m_h$  value can be maximized at loop level for a specific value of  $X_t/M_{\text{SUSY}}$ . For fixed  $X_t$ , the value of  $m_h$  can change by several GeV by varying  $M_{\text{SUSY}}$  within a few TeV or by varying  $m_t$  within its experimental uncertainty, as well as by varying SUSY particle parameters that enter only beyond the one-loop order. Moreover, in the large  $\tan\beta$  regime light staus and/or sbottoms with sizable mixing, governed by the  $\mu$  parameter, yield negative radiative corrections to the mass of the lightest Higgs boson, and can lower it by several GeV [255]. Allowing for experimental and theoretical uncertainties, one finds that for  $M_{\text{SUSY}} \lesssim 2\text{ TeV}$ , large  $m_A$ ,  $\tan\beta \gg 1$  and for  $X_t \simeq \sqrt{6}M_{\text{SUSY}}$ , the maximal value for the lightest Higgs mass is  $m_h^{\text{max}} = 135\text{ GeV}$  [256–257].

<sup>8</sup> Observe that in the SM sections of this review,  $H$  denotes the SM Higgs, whereas in the sections about SUSY, or extensions of the SM with two Higgs doublets,  $H$  is used for the heaviest CP-even Higgs boson, since this is the standard notation in the literature, and the 125 GeV SM-like light Higgs boson will be denoted by  $h$ .

The newly discovered SM-like Higgs boson, if interpreted as the lightest MSSM Higgs with a mass of about 125 GeV, provides information on the possible MSSM parameter space. In particular a sizable mixing in the stop sector is required ( $|X_t/M_{\text{SUSY}}| \geq 1.5$ ) for values of  $M_{\text{SUSY}} \simeq M_Q \simeq M_U \simeq M_D \simeq 1$  to a few TeV [255–269]. See for example Fig. 11.23. On the other hand, considering the third generation soft SUSY breaking parameters as independent inputs,  $M_Q \neq M_U \neq M_D$ , it follows that  $m_h \simeq 125\text{ GeV}$  can be obtained for one stop that is as light as can be experimentally allowed [244] and the other one with a mass of the order of the stop mixing parameter. Considering both stops significantly above a few TeV, by varying/lowering the values of  $X_t$  and  $\tan\beta$ , the impact of higher loops in the computation of the Higgs mass becomes relevant [270–272]. For a given CP-odd Higgs mass  $m_A$ , the masses of the other two Higgs bosons,  $H$  and  $H^\pm$ , also receive radiative corrections. For a more detailed discussion of the effect of radiative corrections on the heavy Higgs masses see for example Refs. [40] and [251].



**Figure 11.23:** Values of the SUSY mass scale  $M_{\text{SUSY}} = M_S$  versus the stop mixing parameter normalized by the SUSY mass scale  $X_t/M_{\text{SUSY}}$ , for fixed  $\tan\beta = 20$ ,  $\mu = 200\text{ GeV}$  and  $M_A = A_t = A_b = A_\tau = M_{\text{SUSY}}$ . The solid black line corresponds to  $M_h = 125\text{ GeV}$  while in the grey band  $M_h$  varies by  $\pm 1\text{ GeV}$ . The red dotted lines are iso-values of the stop mass. This figure is based on Ref. [268].

The phenomenology of the Higgs sector depends on the couplings of the Higgs bosons to gauge bosons and fermions. The couplings of the two CP-even Higgs bosons to  $W$  and  $Z$  bosons are given in terms of the angles  $\alpha$  and  $\beta$

$$g_{hVV} = g_V m_V \sin(\beta - \alpha), \quad g_{HVV} = g_V m_V \cos(\beta - \alpha), \quad (11.26)$$

where  $g_V \equiv 2m_V/v$ , for  $V = W^\pm$  or  $Z$  ( $g_V m_V$  is the SM  $hVV$  coupling). There are no tree-level couplings of  $A$  or  $H^\pm$  to  $VV$ . The couplings of the  $Z$  boson to two neutral Higgs bosons, which must have opposite CP-quantum numbers, are given by  $g_{\phi AZ}(p_\phi - p_A)$ , where  $\phi = H$  or  $h$ , the momenta  $p_\phi$  and  $p_A$  point into the vertex, and

$$g_{hAZ} = g_Z \cos(\beta - \alpha)/2, \quad g_{HAZ} = -g_Z \sin(\beta - \alpha)/2. \quad (11.27)$$

Charged Higgs- $W$  boson couplings to neutral Higgs bosons and four-point couplings of vector bosons and Higgs bosons can be found in Ref. [12].



The tree-level Higgs couplings to fermions obey the following property: the neutral components of one Higgs doublet,  $\Phi_1$ , couple exclusively to down-type fermion pairs while the neutral components of the other doublet,  $\Phi_2$ , couple exclusively to up-type fermion pairs [12]. This Higgs-fermion coupling structure defines the Type-II 2HDM [273]. In the MSSM, fermion masses are generated when both neutral Higgs components acquire vacuum expectation values, and the relations between Yukawa couplings and fermion masses are (in third-generation notation)

$$h_{b,\tau} = \sqrt{2} m_{b,\tau} / (v \cos \beta), \quad h_t = \sqrt{2} m_t / (v \sin \beta). \quad (11.28)$$

The couplings of the neutral Higgs bosons to  $f\bar{f}$  relative to the SM value,  $g_{f\bar{f}}/2M_W$ , are given by

$$\begin{aligned} h\bar{b}b &: -\sin \alpha / \cos \beta & h\bar{t}t &: \cos \alpha / \sin \beta, \\ H\bar{b}b &: \cos \alpha / \cos \beta & H\bar{t}t &: \sin \alpha / \sin \beta, \\ A\bar{b}b &: \gamma_5 \tan \beta & A\bar{t}t &: \gamma_5 \cot \beta. \end{aligned} \quad (11.29)$$

In each relation above, the factor listed for  $\bar{b}b$  also pertains to  $\tau^+\tau^-$ . The charged Higgs boson couplings to fermion pairs are given by

$$\begin{aligned} g_{H-\bar{t}b} &= \frac{g}{\sqrt{2}M_W} \left[ m_t \cot \beta \frac{1+\gamma_5}{2} + m_b \tan \beta \frac{1-\gamma_5}{2} \right], \\ g_{H-\tau^+\nu} &= \frac{g}{\sqrt{2}M_W} \left[ m_\tau \tan \beta \frac{1-\gamma_5}{2} \right]. \end{aligned} \quad (11.30)$$

The non-standard neutral Higgs bosons have significantly enhanced couplings to down-type fermions at sizeable  $\tan \beta$ . In the alignment limit, the lightest Higgs boson behaves like the SM one and  $H$ ,  $A$  have  $\tan \beta$  enhanced couplings to down type fermions, and analogous enhanced couplings are in place for the charged Higgs.

Radiative corrections can modify significantly the values of the Higgs boson couplings to fermion pairs and to vector boson pairs. In a first approximation, when radiative corrections to the quartic couplings are computed, the diagonalizing angle  $\alpha$  is shifted from its tree-level value, and hence one may compute a “radiatively-corrected” value for  $\cos(\beta - \alpha)$  [255, 274]. Additional contributions from the one-loop vertex corrections to tree-level Higgs couplings [276–283] alter significantly the Higgs-fermion Yukawa couplings at large  $\tan \beta$ , both in the neutral and charged Higgs sector.

### VII.1.1. MSSM Higgs boson phenomenology

In the MSSM, the mass, CP properties, decay and production properties of one of the neutral Higgs bosons should agree with the LHC Higgs data. Given that present data allows only for moderate departures from the SM predictions, it implies that some degree of alignment is necessary. For sufficiently heavy non-SM-like Higgs bosons, the alignment results from decoupling. If  $m_A$  is below a few hundred GeV, radiative corrections to the angle  $\alpha$  can result in alignment for sizable values of the Higgs mass parameter  $\mu > M_{\text{SUSY}}$ . Such radiative corrections, being proportional to ratios of mass parameters associated to the SUSY particles, do not decouple for heavy SUSY spectra.

The SM-like branching ratios of  $h$  can be modified if decays into supersymmetric particles are kinematically allowed, and, in particular, decays into a pair of the lightest supersymmetric particles - i.e. the lightest neutralinos,  $\tilde{\chi}_1^0$  - can become dominant and would be invisible if R-parity is conserved [284–286]. Moreover, if light superpartners exist that couple to photons and/or gluons, the  $h$  loop-induced coupling to  $gg$  and  $\gamma\gamma$  could deviate sizably from the corresponding SM predictions [255, 287–294].

Given that some degree of alignment is necessary to agree with data, for the heavier Higgs states there are two possibilities to be considered: i) Alignment triggered by decoupling, hence  $m_A \geq$  several hundred GeV: The  $HWW$  and  $HZZ$  couplings are very small. The dominant  $H$ ,  $A$  decay branching ratios strongly depend on  $\tan \beta$ . After incorporating the leading radiative corrections to Higgs couplings, the following decay features are relevant in the MSSM: The decay modes  $H, A \rightarrow b\bar{b}$ ,  $\tau^+\tau^-$  dominate when  $\tan \beta$  is large (this holds even

away from decoupling). For small  $\tan \beta$ , the  $t\bar{t}$  decay mode dominates above its kinematic threshold. For the charged Higgs boson,  $H^\pm \rightarrow t\bar{b}$  dominates. ii) Some degree of alignment without decoupling, hence  $m_A \leq$  a few hundred GeV. The main difference with the previous case is that in the low  $\tan \beta$  regime ( $\tan \beta \leq 5$ ) additional decay channels may be allowed which involve decays into the lightest SM-like Higgs boson. For  $A$  and  $H$ , besides the  $H, A \rightarrow b\bar{b}$ ,  $\tau^+\tau^-$  decay modes, also  $A \rightarrow Zh$ ,  $H \rightarrow hh$  as well as  $H \rightarrow WW/ZZ$  decay modes are available (they are suppressed in the strict alignment limit). For the charged Higgs boson,  $H^\pm \rightarrow \tau^\pm \nu_\tau$  dominates below the  $t\bar{b}$  threshold, and also  $H^\pm \rightarrow W^\pm h$  may be searched for. Both in i) and ii), the heavier Higgs states,  $H$ ,  $A$  and  $H^\pm$ , are roughly mass degenerate (with masses  $\pm 20$  GeV or less apart). In cases i) and ii) the heavy Higgs boson decays into charginos, neutralinos and third-generation squarks and sleptons can be important if they are kinematically allowed [284].

The main production mechanisms for the neutral MSSM Higgs bosons at  $e^+e^-$  colliders are Higgs-strahlung ( $e^+e^- \rightarrow Zh, ZH$ ), vector boson fusion ( $e^+e^- \rightarrow \nu\bar{\nu}h, \nu\bar{\nu}H$ ) – with  $W^+W^-$  fusion about an order of magnitude larger than  $ZZ$  fusion – and  $s$ -channel  $Z$  boson exchange ( $e^+e^- \rightarrow Ah, AH$ ). For the Higgs-strahlung process it is possible to reconstruct the mass and momentum of the Higgs boson recoiling against the particles from the  $Z$  boson decay, and hence sensitive searches for Higgs bosons decaying to invisible final states are possible. The main charged Higgs boson production process at  $e^+e^-$  colliders is via  $s$ -channel  $\gamma$  or  $Z$  boson exchange ( $e^+e^- \rightarrow H^\pm H^\mp$ ). Charged Higgs bosons can also be produced in top quark decays via  $t \rightarrow b + H^\pm$  if  $m_H^\pm < m_t - m_b$  or via the one-loop process  $e^+e^- \rightarrow W^\pm H^\mp$ , which allows the production of a charged Higgs boson with  $m_H^\pm > \sqrt{s}/2$ , even when  $H^+H^-$  production is kinematically forbidden. Other single charged Higgs production mechanisms include  $t\bar{b}H^-/\bar{t}bH^+$  production,  $\tau^+\nu H^-/\tau^-\bar{\nu}H^+$  production, and a variety of processes in which  $H^\pm$  is produced in association with one or two other gauge and/or Higgs bosons. For representative references on production mechanisms for the MSSM Higgs bosons at  $e^+e^-$  see [243].

At hadron colliders, the dominant neutral Higgs production mechanism at moderate values of  $\tan \beta$  is gluon fusion, mediated by loops containing heavy top and bottom quarks and the corresponding supersymmetric partners [243]. The effect of light stops that may contribute to the gluon fusion production can be partially cancelled by mixing effects. Higgs boson radiation off bottom quarks becomes important for large  $\tan \beta$ , where at least two of the three neutral Higgs bosons have enhanced couplings to bottom-type fermions [295, 296]. In the search for non-standard neutral Higgs bosons,  $A$  and  $H$ , the production can be via either of the above channels in the final inclusive ditau mode and via radiation off bottom quarks in the  $4b$ ’s final mode. The total production rates of bottom quarks and  $\tau$  pairs mediated by the production of a CP-odd Higgs boson in the large  $\tan \beta$  regime are approximately given by

$$\begin{aligned} \sigma_{b\bar{b}A} \times \text{BR}(A \rightarrow b\bar{b}) &\simeq \sigma_{b\bar{b}A}^{\text{SM}} \frac{\tan^2 \beta}{(1 + \Delta_b)^2} \frac{9}{(1 + \Delta_b)^2 + 9}, \\ \sigma_{g\bar{g} \rightarrow A, b\bar{b}A} \times \text{BR}(A \rightarrow \tau^+\tau^-) &\simeq \sigma_{g\bar{g} \rightarrow A, b\bar{b}A}^{\text{SM}} \frac{\tan^2 \beta}{(1 + \Delta_b)^2 + 9}, \end{aligned} \quad (11.31)$$

where  $\sigma_{b\bar{b}A}^{\text{SM}}$  and  $\sigma_{g\bar{g} \rightarrow A, b\bar{b}A}^{\text{SM}}$  denote the values of the corresponding SM Higgs boson cross sections for a SM Higgs boson mass equal to  $m_A$ . For high  $\tan \beta$ , the function  $\Delta_b$  includes the dominant effects of the SUSY radiative corrections affecting the relation between the bottom quark mass and the bottom Yukawa coupling [259, 274, 275, 280–282], and it depends strongly on  $\tan \beta$  and on the SUSY mass parameters. The production and decay rates of  $H$ , for  $m_A$  larger  $m_h^{\text{max}}$ , are governed by formulas similar to the ones presented above, and given that  $A$  and  $H$  are nearly degenerate in mass, the total signal cross section is increased by roughly a factor of two. Detailed discussions of the impact of radiative corrections in these search modes are presented in Refs. [259, 297].

The vector boson fusion and Higgs-strahlung production of the CP-even Higgs bosons as well as the associated production of neutral Higgs bosons with top quark pairs have lower production cross sections

by at least an order of magnitude with respect to the dominant ones, depending on the precise region of MSSM parameter space [42]. Higgs pair production of non-standard MSSM Higgs bosons has been studied in Ref. [298].

Charged Higgs bosons can be produced in several different modes at hadron colliders. If  $m_{H^\pm} < m_t - m_b$ , the charged Higgs boson can be produced in decays of the top quark via the decay  $t \rightarrow bH^+$ , which would compete with the SM process  $t \rightarrow bW^+$ . Relevant radiative corrections to  $\text{BR}(t \rightarrow H^+b)$  have been computed in Refs. [299–302]. For values of  $m_{H^\pm}$  near  $m_t$ , width effects are important. If  $m_{H^\pm} > m_t - m_b$ , then charged Higgs boson production occurs mainly through radiation from a third generation quark. Charged Higgs bosons may also be produced singly in association with a top quark via the  $2 \rightarrow 3$  partonic processes  $pp \rightarrow H^+tb + X$  and  $pp \rightarrow H^-tb + X$ ; via associated production with  $W^\pm$  bosons through  $b\bar{b}$  annihilation and  $gg$  fusion [303]; and in pairs via  $q\bar{q}$  annihilation [304]. The inclusive  $H^+H^-$  cross section is less than the cross section for single charged Higgs associated production [304, 305]. For a more extensive discussion of charged Higgs boson production at LHC see Refs. [1142, 306].

The additional Higgs bosons are sought for mainly via the channels

$$\begin{aligned} pp &\rightarrow A/H \rightarrow \tau^+\tau^- \text{ (inclusive),} \\ b\bar{b}A/H, A/H &\rightarrow \tau^+\tau^- \text{ (with } b\text{-tag),} \\ b\bar{b}A/H, A/H &\rightarrow b\bar{b} \text{ (with } b\text{-tag),} \\ pp &\rightarrow t\bar{t} \rightarrow H^\pm W^\mp b\bar{b}, \quad H^\pm \rightarrow \tau\nu_\tau, \\ gb &\rightarrow H^-t \text{ or } g\bar{b} \rightarrow H^+\bar{t}, \quad H^\pm \rightarrow \tau\nu_\tau. \end{aligned} \quad (11.32)$$

After the Higgs boson discovery, updated MSSM benchmarks scenarios have been defined, that highlight interesting conditions for MSSM Higgs searches [44, 258]. They include: i) a moderate mixing scenario in which the light CP-even Higgs boson can be interpreted as the newly discovered state in most of the  $m_A - \tan\beta$  plane; ii) a light stop scenario with stop masses in the few to several hundred GeV range that can affect gluon fusion Higgs production; and iii) a tau-phobic scenario that exhibits variations of  $\text{BR}(h \rightarrow b\bar{b})$  and  $\text{BR}(h \rightarrow \tau^+\tau^-)$  with respect to their SM values. In the above benchmarks it is also possible to have decays of  $H \rightarrow hh$  in regions of moderate  $m_A$  and moderate  $\tan\beta$  as far as one is away from precise alignment. Also for the previous benchmarks, the LHC reach in the traditional  $A/H \rightarrow \tau^+\tau^-$  search channel varies depending on the values of  $\mu$  and  $M_2$ , that may enable the  $A/H$  decays into electroweakinos. Lastly, varying the parameter  $\mu$  in both sign and magnitude induces relevant variations in the possible discovery reach through the  $4b$ 's channel, and to a lesser extent through the inclusive ditau channel.

An alternative approach to reduce the large number of parameters relevant to the Higgs sector is to consider that, in the Higgs basis, the only important radiative corrections are those affecting the Higgs mass [307]. This approximation is called hMSSM and works well in large regions of parameter space but it breaks down for sizable values of  $\mu$  and  $A_t$ , and moderate values of  $\tan\beta$ , for which the radiative corrections to the mixing between the two CP even eigenstates become relevant. The effect of such radiative corrections is to allow for alignment for small to intermediate values of  $\tan\beta$ , independent of the specific value of  $m_A$  [308]. In addition, the hMSSM assumption that the right value of the Higgs mass may be obtained for all values of  $m_A$  and  $\tan\beta$  is in conflict with the MSSM predictions for the Higgs mass for small values of  $m_A$  and  $\tan\beta \simeq \mathcal{O}(1)$ .

Future precision measurements of the Higgs boson couplings to fermions and gauge bosons together with information on heavy Higgs searches will provide powerful information on the SUSY parameter space [309, 308–311]. If no other new states beyond the current Higgs candidate are discovered at the LHC, it becomes mandatory to understand what would be the required precision of the Higgs rate measurements to distinguish the MSSM from the SM.

Improvements in our understanding of  $B$ -physics observables put indirect constraints on additional Higgs bosons in mass ranges that would be accessible in direct LHC searches. In particular,  $\text{BR}(B_s \rightarrow \mu^+\mu^-)$ ,  $\text{BR}(b \rightarrow s\gamma)$ , and  $\text{BR}(B_u \rightarrow \tau\nu)$  play an important

**Table 11.16:** Symmetries associated to various models with singlet extensions, the corresponding terms in the superpotential that only involve Higgs and singlet fields, and the number of neutral states in the Higgs sector for the case of CP conservation.

Model	MSSM	NMSSM	nMSSM	UMSSM
Symmetry	-	$Z_3$	$Z_5^R, Z_7^R$	$U(1)'$
Superpotential	$\mu\Phi_2 \cdot \Phi_1$	$\lambda_S S\Phi_2 \cdot \Phi_1 + \frac{\kappa}{3}S^3$	$\lambda_S S\Phi_2 \cdot \Phi_1 + t_F S$	$\lambda_S S\Phi_2 \cdot \Phi_1$
$H_i^0$	2	3	3	3
$A_i^0$	1	2	2	1

role within minimal flavor-violating (MFV) models [312], in which flavor effects proportional to the CKM matrix elements are induced, as in the SM. For example, see references in [243]. The supersymmetric contributions to these observables come both at the tree and loop level, and have a different parametric dependence, but share the property that they become significant for large values of  $\tan\beta$ , which is also the regime in which searches for non-standard MSSM Higgs bosons at hadron colliders are the most powerful.

## VII.2. Higgs bosons in singlet extensions of the MSSM

In the MSSM, the Higgs mass parameter  $\mu$  is a supersymmetric parameter, and as such, it should naturally be of order  $M_{GUT}$  or  $M_{Planck}$ . The fact that phenomenologically it is required that  $\mu$  be at the electroweak/TeV scale is known as the  $\mu$  problem [313]. Supersymmetric models with additional singlets can provide a solution to the  $\mu$  problem, by promoting the  $\mu$  parameter to a dynamical singlet superfield  $S$  that only interacts with the MSSM Higgs doublets through a coupling  $\lambda_S$  at the level of the superpotential. An effective  $\mu$  is generated when the real scalar component of  $S$  acquires a vacuum expectation value  $\langle S \rangle$

$$\mu_{eff} = \lambda_S \langle S \rangle. \quad (11.33)$$

After the minimization of the Higgs potential the vacuum state relates the vacuum expectation values of the three CP-even neutral scalars,  $\phi_1^0, \phi_2^0$  and  $S$ , to their soft supersymmetry breaking masses, hence, one expects that these VEVs should all be of order  $M_{SUSY}$  and therefore the  $\mu$  problem is solved.

The solution of the  $\mu$  problem through the addition of a singlet superfield to the MSSM comes along with the existence of an extra global  $U(1)$  symmetry, known as the Peccei–Quinn (PQ) symmetry [314]. This PQ symmetry is broken explicitly in realistic models. For that purpose one can consider a discrete  $Z_3$  symmetry that allows the existence of a PQ odd  $S^3$  term in the superpotential. This model extension has been called the next-to-minimal supersymmetric SM (NMSSM) (the NMSSM was first introduced in Ref. [315], see the 2014 edition of this review for an extensive list of subsequent references). It is known however that discrete symmetries may come along with the existence of domain wall structures that imply that our universe would consist of disconnected domains with different ground states, creating unacceptably large anisotropies in the cosmic microwave background [316]. To avoid the problem of domain walls one can consider the existence of non-renormalizable operators that would lead to the preferred vacuum state. However, the same operators in turn may generate quadratically divergent tadpole contributions [243] that could shift the VEV of  $S$  to be much larger, order  $M_{GUT}$ , and ruin the singlet solution to the  $\mu$  problem. To cure the problem of destabilizing tadpoles, discrete R-symmetries have been proposed that ensure that tadpoles would only appear at very high order loops and be safely suppressed. Depending on the symmetries imposed on the theory, different models with singlet extensions of the MSSM (xMSSM) have been proposed. In Table 11.16 we show the most studied examples: the NMSSM, the Nearly-Minimal Supersymmetric SM (nMSSM) [317], and the  $U(1)'$ -extended MSSM (UMSSM) [318], specifying the new parameters appearing in the superpotential and the respective symmetries. A Secluded  $U(1)'$ -extended MSSM (sMSSM) [319] contains three singlets in addition to the standard UMSSM Higgs singlet; this model is equivalent to the nMSSM in the limit that the additional singlet VEV's are large, and the trilinear singlet coupling,  $\lambda_S$ , is small [320].

Based on the extended models defined in Table 11.16, we write the most generic supersymmetric and soft supersymmetry breaking scalar potentials for the three scalar fields:  $\Phi_1$ ,  $\Phi_2$  and  $S$ :

$$V_{xMSSM} = \left| \lambda_S \Phi_2 \cdot \Phi_1 + t_F + \kappa S^2 \right|^2 + |\lambda_S S|^2 (|\Phi_1|^2 + |\Phi_2|^2) + \frac{g'^2 + g^2}{8} (|\Phi_1|^2 - |\Phi_2|^2)^2 + \frac{g^2}{2} (|\Phi_1|^2 |\Phi_2|^2 - |\Phi_2 \cdot \Phi_1|^2) + \frac{g_1'^2}{2} (Q_{\Phi_1} |\Phi_1|^2 + Q_{\Phi_2} |\Phi_2|^2 + Q_S |S|^2)^2 \quad (11.34)$$

$$V_{\text{soft}} = m_{H_1}^2 |\Phi_1|^2 + m_{H_2}^2 |\Phi_2|^2 + m_s^2 |S|^2 + \left( A_s \lambda_S S H_u \cdot H_d + \frac{\kappa}{3} A_\kappa S^3 + t_S S + h.c. \right). \quad (11.35)$$

where  $\Phi_2 \cdot \Phi_1 = \epsilon_{ij} \Phi_2^i \Phi_1^j$  and the couplings  $g'$ ,  $g$ , and  $g_1'$  are associated to the  $U(1)_Y$ ,  $SU(2)_L$ , and  $U(1)'$  gauge symmetries, respectively.  $t_F$  and  $t_S$  are supersymmetric and SUSY breaking tadpole terms, respectively,  $m_s$  is a SUSY breaking mass term for the scalar component of the field  $S$ , and  $A_s$  and  $A_\kappa$  are the trilinear soft SUSY breaking mass parameters associated with the new terms  $\lambda_S S \Phi_2 \cdot \Phi_1$  and  $\kappa S^3/3$  in the superpotential, with the B-term of the MSSM expressed as  $B\mu \equiv A_s \mu_{\text{eff}}$ . In particular,  $\kappa$  and  $A_\kappa$  are the parameters for the NMSSM model, while  $t_F$  and  $t_S$  are those of the nMSSM. The UMSSM depends on the new coupling  $g_1'$  as well as on the  $U(1)'$  charges of the Higgs fields,  $Q_{\Phi_1}$ ,  $Q_{\Phi_2}$  and  $Q_S$ , that are free parameters with the restriction that they have to add to zero for the superpotential  $\lambda_S S \Phi_2 \Phi_1$  to be gauge invariant. In a given  $U(1)'$  construction the charges are specified. The addition of the singlet scalar field(s) imply that additional CP-even and CP-odd Higgs bosons will appear in the spectra, whereas the charged Higgs sector remains the same as in the MSSM given that the number of Higgs doublets remains unchanged. The mixing with the extra scalar  $S$  alters the masses and properties of the physical Higgs bosons, that in general can differ significantly from the SM or the MSSM. A detailed discussion of typical mass spectra and decay properties in these models can be found for example in Refs. [321, 320]. Moreover, these models have extra neutralinos and in some cases extra neutral gauge bosons,  $Z'$ . The extra gauge boson sector is constrained by experimental data through direct  $Z'$  searches as well as the  $Z - Z'$  mixing angle  $\alpha_{ZZ'}$  constrained to be less than  $\mathcal{O}(10^{-3})$  by precision electroweak data.

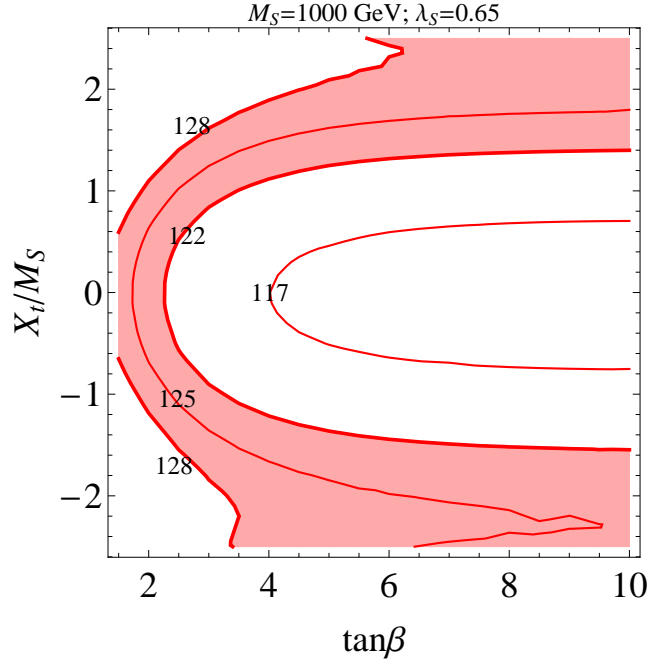
In singlet extensions of the MSSM the lightest CP-even Higgs mass at tree level,  $m_{H_1}^{\text{tree}}$  receives a contribution from the singlet scalar that renders it larger than the MSSM value, in particular for small values of  $\tan\beta$ . The tree level upper bound reads<sup>9</sup>

$$m_{H_1}^{\text{tree}} \leq M_Z^2 \cos^2 2\beta + \frac{1}{2} \lambda_S^2 v^2 \sin^2 2\beta. \quad (11.36)$$

At the one-loop level, the top and stop loops (as well as sbottom and stau loops for large  $\tan\beta$ ) are the dominant contributions, that are common to the MSSM and to all the singlet extensions. Gauge couplings in the MSSM are small compared to the top quark Yukawa coupling, hence the one-loop gauge contributions are negligible. Corrections exclusive to the NMSSM and the nMSSM enter only at the two loop level. Therefore, there are no significant model-dependent contributions at one loop order, and as a result, for large  $\tan\beta$  the lightest CP-even Higgs mass does not differ in any significant way from the MSSM one. In the decoupling limit, a value of the lightest SM Higgs mass of about 125 GeV is achievable in all these MSSM extensions, and this remains the case even after higher order corrections are implemented.

A singlet extended supersymmetric Higgs sector opens new avenues for discovery. Since the singlet pseudoscalar particle may be identified as the pseudo-Goldstone boson of a spontaneously broken Peccei–Quinn symmetry, it may become naturally light [322, 323]. Generally, there is mixing of the singlet sector with the MSSM Higgs sector, and for a sufficiently light, singlet dominated scalar or pseudoscalar,  $h_S$  or

<sup>9</sup> Additional gauge interactions in the UMSSM contribute to this increase with a term of  $\mathcal{O}(g_1'^2 v^2 (Q_{\Phi_2}^2 \cos^2 \beta + Q_{\Phi_1}^2 \sin^2 \beta))$ .



**Figure 11.24:** Values of the stop mixing parameter normalized to the SUSY mass scale  $X_t/M_{\text{SUSY}}$ , as a function of  $\tan\beta$ , for  $M_{\text{SUSY}} \equiv M_S = 1000$  GeV,  $\lambda_S = 0.65$ , and contours of constant values of the Higgs mass  $m_h = 125 \pm 3$  GeV shaded in red [327].

$A_S$ , respectively, the SM-like Higgs boson  $h$  may decay to pairs of  $h_S$  or  $A_S$ . The light scalar and/or pseudoscalar may subsequently decay to  $\tau\tau$  or  $b\bar{b}$  pairs. Such cascade decays are more difficult to detect than standard searches due to the potentially soft decay products. There is also a rich phenomenology for the decays of the heavy CP-even and CP-odd doublets,  $A$  and  $H$  into two lighter Higgs bosons such as  $H \rightarrow hh_S$ ,  $hh$ ,  $h_S h_S$  or  $A \rightarrow A_S h_S$ ,  $A_S h$  as well as into a light Higgs boson and a gauge boson:  $H \rightarrow A_S Z$ ;  $A \rightarrow h_S Z$ ,  $hZ$ . If kinematically allowed the heavy Higgs bosons decay into  $t\bar{t}$ . If the singlet dominated scalar or pseudoscalar are somewhat heavier, the decays  $h_S \rightarrow WW$  or  $A_S \rightarrow h_S Z$  will be allowed.

In addition, the light singlet scenario in the NMSSM or nMSSM is typically associated with a light singlino-dominated neutralino. The recently discovered SM-like Higgs boson can then decay to pairs of this neutralino [324, 320], opening an invisible decay mode that is not excluded by present data. All of the Higgs bosons can decay into electroweakinos depending on kinematics and the singlino or higgsino composition of the electroweakinos.

In models with extended singlets, at low  $\tan\beta$  it is possible to trade the requirement of a large stop mixing by a sizeable trilinear Higgs-singlet Higgs coupling  $\lambda_S$ , rendering more freedom on the requirements for gluon fusion production. As in the MSSM, mixing in the Higgs sector - additionally triggered by the extra new parameter  $\lambda_S$  - can produce variations in the Higgs- $b\bar{b}$  and Higgs- $\tau^+ \tau^-$  couplings that can alter the Higgs to  $ZZ/WW$  and diphoton rates. Light charginos at low  $\tan\beta$  can independently contribute to enhance the diphoton rate, without altering any other of the Higgs decay rates [290, 325].

There is much activity in exploring the NMSSM phenomenology in the light of the 125 GeV Higgs boson [326], as well as in defining benchmark scenarios with new topologies including Higgs decay chains [291]. An analytic understanding of the alignment condition in the NMSSM is presented in Ref. [292]. The NMSSM with a Higgs boson of mass 125 GeV can be compatible with stop masses of order of the electroweak scale, thereby reducing the degree of fine tuning necessary to achieve electroweak symmetry breaking (see Fig. 11.24). The alignment conditions point toward a more natural region of parameter space for electroweak symmetry breaking, while allowing for perturbativity of the theory up to the Planck scale and yielding a rich and interesting Higgs boson phenomenology at the LHC.

### VII.3. Supersymmetry with extended gauge sectors

In the MSSM, the tree-level value of the lightest CP-even Higgs mass originates from the D-term dependence of the scalar potential that comes from the supersymmetric kinetic terms in the Kähler potential. The D-terms lead to tree-level quartic couplings which are governed by the squares of the gauge couplings of the weak interactions, under which the Higgs has non-trivial charges and hence the lightest Higgs mass is bounded to be smaller than  $M_Z$ . If new gauge interactions were present at the TeV scale, and the Higgs bosons would have non-trivial charges under them, there would be new D-term contributions that would lead to an enhancement of the tree-level Higgs mass value. Since the low energy gauge interactions reduce to the known  $SU(3)_c \times SU(2)_L \times U(1)_Y$  ones, in order for this mechanism to work, the extended gauge and Higgs sectors should be integrated out in a non-supersymmetric way. This means that there must be supersymmetry breaking terms that are of the order of, or larger than, the new gauge boson masses. The tree-level quartic couplings would then be enhanced through their dependence on the square of the gauge couplings of the extended Higgs sector. This effect will be suppressed when the heavy gauge boson masses are larger than the supersymmetry breaking scale and will acquire its full potential only for large values of this scale.

One of the simplest possibilities is to extend the weak interactions to a  $SU(2)_1 \times SU(2)_2$  sector, such that the known weak interactions are obtained after the spontaneous breaking of these groups to  $SU(2)_L$  [328]. This may be achieved by introducing a bi-doublet  $\Sigma$  under the two  $SU(2)$  gauge groups, which acquires a non-trivial vacuum expectation value  $u$  in the diagonal direction. The heavy gauge boson masses are therefore given by  $M_{W'}^2 = (g_1^2 + g_2^2)u^2/2$ , and the weak coupling  $g^2 = g_1^2 g_2^2 / (g_1^2 + g_2^2)$ . To obtain a new tree-level contribution to the Higgs potential, the Higgs bosons must be charged under the new gauge interactions. One possibility is to assume that the third generation quarks and leptons as well as the Higgs doublets have charges under the  $SU(2)_1$  group, while the second and first generations have charges under  $SU(2)_2$ . This provides a natural explanation of the largeness of the third generation couplings compared to the first and second generation ones.

Under the above conditions, the D-term contributions to the neutral Higgs effective potential are given by

$$V_D = \frac{g^2 \Delta + g'^2}{8} (|H_2^0|^2 - |H_1^0|^2)^2 \quad (11.37)$$

with

$$\Delta = \left(1 + \frac{4m_\Sigma^2}{g_2^2 u^2}\right) \left(1 + \frac{4m_\Sigma^2}{(g_1^2 + g_2^2)u^2}\right)^{-1}, \quad (11.38)$$

where  $m_\Sigma$  is the supersymmetry breaking term associated with the bi-doublet  $\Sigma$ . It is easy to see that while the MSSM D-term is recovered when  $m_\Sigma \rightarrow 0$ , it is replaced by the  $SU(2)_1 \times U(1)_Y$  D-term when  $m_\Sigma$  becomes much larger than  $M_{W'}$ . The tree-level mass now reads

$$m_h^2|_{\text{tree}} = \frac{g^2 \Delta + g'^2}{4} v^2 \cos^2 2\beta, \quad (11.39)$$

and reduces to the MSSM value,  $M_Z^2 \cos^2 2\beta$ , for  $\Delta = 1$ .

Assuming  $g_1 \simeq g_2$ , values of  $g_{1,2}$  of order one are necessary to obtain the proper value of the weak gauge coupling. In addition, if values of  $m_\Sigma$  of order  $M_{W'}$  are assumed, enhancements of order 50 percent of the MSSM D-term contribution to the Higgs mass may be obtained. Such enhancements are sufficient to obtain the measured Higgs mass value without the need for very heavy stops or large stop mixing parameters.

The gauge extension described above leads to new, heavy gauge and Higgs bosons, as well as new neutralinos and charginos. Constraints from precision measurements put bounds of the order of a few TeV on the mass of these gauge bosons, which may be probed at the higher energy run of the LHC collider. If the new gaugino supersymmetry breaking masses are smaller than the gauge boson masses, the new electroweakinos will have masses of the order of a few TeV and therefore the weak scale phenomenology reduces to the MSSM one.

Similar gauge extensions, including also new abelian gauge groups have been considered, for instance, in Ref. [329].

Gauge extensions of the MSSM can also lead to an enhancement of the Higgs mass value by modifying the renormalization group evolution of the Higgs quartic coupling to low energies. In the MSSM, the evolution of the quartic coupling is governed by the top-quark Yukawa interactions and depends on the fourth power of the top-quark Yukawa coupling. The neutralino and chargino contributions, which depend on the fourth power of the weak gauge couplings, are small due to the smallness of these couplings. Depending on the values of the soft supersymmetry breaking parameters in the gaugino and Higgsino sectors, the  $SU(2)_1$  gauginos may become light, with masses of the order of the weak scale. Since the  $SU(2)_1$  coupling may be significantly larger than the  $SU(2)_L$  one, for small values of the Higgsino mass parameter  $\mu$ , the associated charginos and neutralinos may modify the evolution of the quartic coupling in a significant way [330]. This may lead to a significant increase of the lightest CP-even Higgs mass, even for small values of  $\tan \beta \simeq 1$  for which the D-term contributions become small. In addition, under these conditions, light charginos may lead to a significant modification of the Higgs diphoton decay rate, which may be as large as 50% of the SM [330–334].

### VII.4. Effects of CP violation

SUSY scenarios with CP-violation (CPV) phases are theoretically appealing, since additional CPV beyond that observed in the  $K$ ,  $D$ , and  $B$  meson systems is required to explain the observed cosmic matter-antimatter asymmetry. In the MSSM, CP-violation effects in the Higgs sector appear at the quantum level, while in singlet extensions of the MSSM CP-violation effects can already be effective at tree level. In general, CP-violation effects in the Higgs sector have significant constraints from electric dipole moments data [243].

In the MSSM, the gaugino mass parameters ( $M_{1,2,3}$ ), the Higgsino mass parameter,  $\mu$ , the bilinear Higgs squared-mass parameter,  $m_{12}^2$ , and the trilinear couplings of the squark and slepton fields to the Higgs fields,  $A_f$ , may carry non-trivial phases. The two parameter combinations  $\arg[\mu A_f (m_{12}^2)^*]$  and  $\arg[\mu M_i (m_{12}^2)^*]$  are invariant under phase redefinitions of the MSSM fields [335, 261]. Therefore, if one of these quantities is non-zero, there would be new sources of CP-violation, which affects the Higgs sector through radiative corrections [260, 261, 336–340]. The mixing of the neutral CP-odd and CP-even Higgs boson states is no longer forbidden. Hence,  $m_A$  is no longer a physical parameter. However, the charged Higgs boson mass  $m_{H^\pm}$  is still physical and can be used as an input for the computation of the neutral Higgs spectrum of the theory. For large values of  $m_{H^\pm}$ , corresponding to the decoupling limit, the properties of the lightest neutral Higgs boson state approach those of the SM Higgs boson. In particular, the upper bound on the lightest neutral Higgs boson mass, takes the same value as in the CP-conserving case [261]. Nevertheless, there still can be significant mixing between the two heavier neutral mass eigenstates. For a detailed study of the Higgs boson mass spectrum and parametric dependence of the associated radiative corrections, see Refs. [336, 339].

Major variations to the MSSM Higgs phenomenology occur in the presence of explicit CPV phases. In the CPV case, vector boson pairs couple to all three neutral Higgs boson mass eigenstates,  $H_i$  ( $i = 1, 2, 3$ ), with couplings

$$g_{H_i V V} = \cos \beta \mathcal{O}_{1i} + \sin \beta \mathcal{O}_{2i}, \quad (11.40)$$

$$g_{H_i H_j Z} = \mathcal{O}_{3i} (\cos \beta \mathcal{O}_{2j} - \sin \beta \mathcal{O}_{1j}) - \mathcal{O}_{3j} (\cos \beta \mathcal{O}_{2i} - \sin \beta \mathcal{O}_{1i}), \quad (11.41)$$

where the  $g_{H_i V V}$  couplings are normalized to the analogous SM coupling and the  $g_{H_i H_j Z}$  have been normalized to  $g_Z^{\text{SM}}/2$ . The orthogonal matrix  $\mathcal{O}_{ij}$  is relating the weak eigenstates to the mass eigenstates. It has non-zero off-diagonal entries mixing the CP-even and CP-odd components of the weak eigenstates. The above couplings obey the relations

$$\sum_{i=1}^3 g_{H_i Z Z}^2 = 1 \quad \text{and} \quad g_{H_k Z Z} = \varepsilon_{ijk} g_{H_i H_j Z}, \quad (11.42)$$

where  $\varepsilon_{ijk}$  is the Levi-Civita symbol. Moreover, CPV phases imply that all neutral Higgs bosons can couple to both scalar and pseudoscalar fermion bilinear densities. The couplings of the mass eigenstates  $H_i$  to fermions depend on the loop-corrected fermion Yukawa couplings (similarly to the CPC case), on  $\tan\beta$  and on the  $\mathcal{O}_{ji}$  [336, 341].

The production processes of neutral MSSM Higgs bosons in the CPV scenario are similar to those in the CPC scenario. Regarding the decay properties, the lightest mass eigenstate,  $H_1$ , predominantly decays to  $b\bar{b}$  if kinematically allowed, with a smaller fraction decaying to  $\tau^+\tau^-$ , similar to the CPC case. If kinematically allowed, a SM-like neutral Higgs boson,  $H_2$  or  $H_3$  can decay predominantly to  $H_1H_1$  leading to many new interesting signals both at lepton and hadron colliders; otherwise it will decay preferentially to  $b\bar{b}$ .

The discovery of a 125 GeV Higgs boson has put strong constraints on the realization of the CPV scenario within the MSSM. This is partly due to the fact that the observed Higgs rates are close to the SM values, and a large CP-violating component would necessarily induce a large variation in the rate of the SM-like Higgs decay into the weak gauge bosons  $W^\pm$  and  $Z$ . The measured Higgs mass imposes additional constraints on the realization of this scenario. Once all effects are considered, the CP-odd Higgs  $A$  component of the lightest Higgs tends to be smaller than about 10%. This restriction can be alleviated in the NMSSM or more general two Higgs doublet models. CP-violating effects can still be significant in the heavy Higgs sector. For instance, the Higgs bosons  $H_2$  and  $H_3$  may be admixtures of CP-even and CP-odd scalars, and therefore both may be able to decay into pairs of weak gauge bosons. The observation of such decays would be a clear signal of CP-violation. In the MSSM the proximity of the masses of  $H_2$  and  $H_3$  makes the measurement of such effect quite challenging, but in generic two Higgs doublet models, the mass splitting between the two heavy mass eigenstates may become larger, facilitating the detection of CP-violating effects at collider experiments. For a recent studies see for example [342].

### VII.5. Non-supersymmetric extensions of the Higgs sector

There are many ways to extend the minimal Higgs sector of the SM. In the preceding sections the phenomenology of SUSY Higgs sectors is considered, which at tree level implies a constrained type-II 2HDM (with restrictions on the Higgs boson masses and couplings). In the following discussion, more generic 2HDM's [12, 273, 343, 344] are presented. These models are theoretically less compelling since they do not provide an explanation for the SM Higgs naturalness problem, but can lead to different patterns of Higgs-fermion couplings, hence, to different phenomenology. It is also possible to consider models with a SM Higgs boson and one or more additional scalar  $SU(2)$  doublets that acquire no VEV and hence play no role in the EWSB mechanism. These models are dubbed Inert Higgs Doublet Models (IHD) [345]. Without a VEV associated to it, a Higgs boson from an inert doublet has no tree-level coupling to gauge bosons and hence cannot decay into a pair of them. And imposing a  $Z_2$  symmetry that prevents them from coupling to the fermions, it follows that, if the lightest inert Higgs boson is neutral, it becomes a good dark matter candidate with interesting associated collider signals. Recent studies of IHD models in the light of a 125 GeV Higgs have been performed [346], showing that there can be non-negligible enhancement or suppression of Higgs to diphotons or Higgs to  $Z\gamma$ . This may be due to the presence of a light charged Higgs, as light as 100 GeV, that is not in conflict with collider or flavor constraints, because it has no couplings to fermions. It is interesting to study the interplay between collider and direct dark matter detection signals in these models.

Other extensions of the Higgs sector can include [321, 347] multiple copies of  $SU(2)_L$  doublets, additional Higgs singlets [348], triplets or more complicated combinations of Higgs multiplets. It is also possible to enlarge the gauge symmetry beyond  $SU(2)_L \times U(1)_Y$  along with the necessary Higgs structure to generate gauge boson and fermion masses. There are two main experimental constraints on these extensions: (i) precision measurements which constrain  $\rho = m_W^2/(m_Z^2 \cos^2\theta_W)$  to be very close to 1 and (ii) flavor changing neutral current (FCNC) effects. In electroweak models based on the SM gauge group, the

tree-level value of  $\rho$  is determined by the Higgs multiplet structure. By suitable choices for the hypercharges, and in some cases the mass splitting between the charged and neutral Higgs sector or the vacuum expectation values of the Higgs fields, it is possible to obtain a richer combination of singlets, doublets, triplets and higher multiplets compatible with precision measurements [349]. Concerning the constraints coming from FCNC effects, the Glashow–Weinberg (GW) criterion [350] states that, in the presence of multiple Higgs doublets the tree-level FCNC's mediated by neutral Higgs bosons will be absent if all fermions of a given electric charge couple to no more than one Higgs doublet. An alternative way of suppressing FCNC in a two Higgs doublet model has been considered in Ref. [351], where it is shown that it is possible to have tree level FCNC completely fixed by the CKM matrix, as a result of an abelian symmetry.

#### VII.5.1. Two-Higgs-doublet models

Supersymmetry demands the existence of two Higgs doublets such that one doublet couples to up-type quarks and the other to down-type quarks and charged leptons. This Higgs-fermion coupling structure is the one identified as type-II 2HDM [273] and assures that masses for both up and down-type quarks can be generated in a supersymmetric and gauge invariant way. Two Higgs doublet models [343], however, can have a more diverse Higgs-fermion coupling structure and can be viewed as a simple extension of the SM to realize the spontaneous breakdown of  $SU(2)_L \times U(1)_Y$  to  $U(1)_{\text{em}}$ . Quite generally, if the two Higgs doublets contain opposite hypercharges, the scalar potential will contain mixing mass parameters of the kind  $m_{12}^2 \Phi_1^T i\sigma_2 \Phi_2 + h.c..$  In the presence of such terms, both Higgs doublets will acquire vacuum expectation values,  $v_1/\sqrt{2}$  and  $v_2/\sqrt{2}$ , respectively, and the gauge boson masses will keep their SM expressions with the Higgs vacuum expectation value  $v$  replaced by  $v = \sqrt{v_1^2 + v_2^2}$ . Apart from the mass terms, the most generic renormalizable and gauge invariant scalar potential contains seven quartic couplings, which are defined in Eq. (11.24).

Considering two doublets with hypercharges, with  $Y_{\Phi_1} = -1$  and  $Y_{\Phi_2} = 1$  as in Eq. (11.23), and the most general, renormalizable Higgs potential will be given by Eq. (11.24). Just as in the MSSM case, after electroweak symmetry breaking and in the absence of CP-violation, the physical spectrum contains a pair of charged Higgs bosons  $H^\pm$ , a CP-odd Higgs boson  $A$  and two neutral CP-even Higgs bosons,  $h$  and  $H$ . The angles  $\alpha$  and  $\beta$  diagonalize the CP-even, and the CP-odd and Charged Higgs sectors, respectively.

The complete 2HDM is defined only after considering the interactions of the Higgs fields to fermions. Yukawa couplings of the generic form

$$-h_{ij}^a \bar{\Psi}_L^i H_a \Psi_R^j + h.c. \quad (11.43)$$

may be added to the renormalizable Lagrangian of the theory. Contrary to the SM, the two Higgs doublet structure does not ensure the alignment of the fermion mass terms  $m_{ij} = h_{ij}^a v_a/\sqrt{2}$  with the Yukawa couplings  $h_{ij}^a$ . This implies that quite generally, the neutral Higgs boson will mediate flavor changing interactions between the different mass eigenstates of the fermion fields. Such flavor changing interactions should be suppressed in order to describe properly the Kaon, D and B meson phenomenology. Based on the Glashow–Weinberg criterion, it is clear that the simplest way of avoiding such transitions is to assume the existence of a symmetry that ensures the couplings of the fermions of each given quantum number (up-type and down-type quarks, charged and neutral leptons) to only one of the two Higgs doublets. Different models may be defined depending on which of these fermion fields couple to a given Higgs boson, see Table 11.17. Models of type-I [344] are those in which all SM fermions couple to a single Higgs field. In type-II models [273] down-type quarks and charged leptons couple to a common Higgs field, while the up-type quarks and neutral leptons couple to the other. In models of type-III (lepton-specific) quarks couple to one of the Higgs bosons, while leptons couple to the other. Finally, in models of type-IV (flipped), up-type quarks and charged leptons couple to one of the Higgs fields while down-quarks and neutral leptons couple to the other.

The two Higgs doublet model phenomenology depends strongly on the size of the mixing angle  $\alpha$  and therefore on the quartic couplings.

**Table 11.17:** Higgs boson couplings to up, down and charged lepton-type  $SU(2)_L$  singlet fermions in the four discrete types of 2HDM models that satisfy the Glashow–Weinberg criterion, from Ref. [352].

Model	2HDM I	2HDM II	2HDM III	2HDM IV
u	$\Phi_2$	$\Phi_2$	$\Phi_2$	$\Phi_2$
d	$\Phi_2$	$\Phi_1$	$\Phi_2$	$\Phi_1$
e	$\Phi_2$	$\Phi_1$	$\Phi_1$	$\Phi_2$

For large values of  $m_A$ ,  $\sin \alpha \rightarrow -\cos \beta$ ,  $\cos \alpha \rightarrow \sin \beta$ ,  $\cos(\beta - \alpha) \rightarrow 0$ , and the lightest CP-even Higgs  $h$  behaves as the SM Higgs. The same behavior is obtained if the quartic couplings are such that  $\mathcal{M}_{12}^2 \sin \beta = -(\mathcal{M}_{11}^2 - m_h^2) \cos \beta$ . The latter condition represents a situation in which the coupling of  $h$  to fermions and weak gauge bosons become the same as in the SM, without decoupling the rest of the non-standard scalars and it is of particular interest due to the fact that the recently discovered Higgs boson has SM-like properties. This situation will be referred to as alignment, as in the MSSM case.

In type-II Higgs doublet models, at large values of  $\tan \beta$  and moderate values of  $m_A$ , the non-standard Higgs bosons  $H, A$  and  $H^\pm$  couple strongly to bottom quarks and  $\tau$  leptons. Hence the decay modes of the non-standard Higgs bosons tend to be dominated by b-quark and tau-lepton modes, including top quarks or neutrinos in the case of the charged Higgs. However, for large and negative values of  $\lambda_4$ , the charged Higgs boson mass may be sufficiently heavy to allow on-shell decays

$$H^\pm \rightarrow W^\pm + (H, A),$$

$$g_{H^\pm W^\mp H, A} \simeq \frac{M_W}{v} \sin(\beta - \alpha)(p_{H^\pm} - p_{H, A}), \quad (11.44)$$

where  $p_{H^\pm}$  and  $p_{H, A}$  are the charged and neutral scalar Higgs momenta pointing into the vertex. On the other hand, for large and positive values of  $\lambda_5$ , the above charged Higgs decay into a  $W^\pm$  and the CP-odd Higgs boson may be allowed, but the heavy Higgs  $H$  may be sufficiently heavy to decay into a CP-odd Higgs boson and an on-shell  $Z$ .

$$H \rightarrow Z + A, \quad g_{HZA} \simeq \frac{M_Z}{v} \sin(\beta - \alpha)(p_H - p_A). \quad (11.45)$$

The decay  $H^\pm \rightarrow W^\pm + H$ , on the other hand may be allowed only if  $\lambda_4 < -\lambda_5$ . The couplings controlling all the above decay modes are proportional to  $\sin(\beta - \alpha)$  and therefore they are unsuppressed in the alignment limit. Moreover, these could still be the dominant decay modes at moderate values of  $\tan \beta$ , offering a way to evade the current bounds obtained assuming a dominant decay into bottom quarks or  $\tau$  leptons.

The quartic couplings are restricted by the condition of stability of the effective potential as well as by the restriction of obtaining the proper value of the lightest CP-even Higgs mass. Close to the alignment limit, the lightest CP-even Higgs mass becomes approximately independent of  $m_A$  and is given by

$$m_h^2 \simeq v^2(\lambda_1 \cos^4 \beta + \lambda_2 \sin^4 \beta + 2\tilde{\lambda}_3 v^2 \cos^2 \beta \sin^2 \beta) + v^2(4\lambda_6 \cos^3 \beta \sin \beta + 4\lambda_7 \sin^3 \beta \cos \beta), \quad (11.46)$$

where  $\tilde{\lambda}_3 = \lambda_3 + \lambda_4 + \lambda_5$ .

The stability conditions imply the positiveness of all masses, as well as the avoidance of run-away solutions to large negative values of the fields in the scalar potential. These conditions imply

$$\begin{aligned} \lambda_1 &\geq 0, \quad \lambda_2 \geq 0, \quad \lambda_3 + \lambda_4 - |\lambda_5| \geq -\sqrt{\lambda_1 \lambda_2}, \\ \lambda_3 &\geq -\sqrt{\lambda_1 \lambda_2}, \quad 2|\lambda_6 + \lambda_7| < \frac{\lambda_1 + \lambda_2}{2} + \tilde{\lambda}_3, \end{aligned} \quad (11.47)$$

where the first four are necessary and sufficient conditions in the case of  $\lambda_6 = \lambda_7 = 0$ , while the last one is a necessary condition in the case all couplings are non-zero. Therefore, to obtain the conditions that allow the decays  $H^\pm \rightarrow W^\pm H, A$  and  $H \rightarrow ZA$ ,  $\lambda_3$  should take

large positive values in order to compensate for the effects of  $\lambda_4$  and  $\lambda_5$ . For recent detailed discussions about 2HDM phenomenology see Refs [309, 254, 347, 353–356, 357].

### VII.5.2. Higgs triplets

Electroweak triplet scalars are the simplest non-doublet extension of the SM that can participate in the spontaneous breakdown of  $SU(2)_L \times U(1)_Y$  to  $U(1)_{\text{em}}$ . Two types of model have been developed in enough detail to make a meaningful comparison to LHC data: the Higgs triplet model (HTM) [358, 359] and the Georgi–Machacek model [360–362].

The Higgs triplet model extends the SM by the addition of a complex  $SU(2)_L$  triplet scalar field  $\Delta$  with hypercharge  $Y = 2$ , and a general gauge-invariant renormalizable potential  $V(\Phi, \Delta)$  for  $\Delta$  and the SM Higgs doublet  $\Phi$ . The components of the triplet field can be parameterized as

$$\Delta = \frac{1}{\sqrt{2}} \begin{pmatrix} \Delta^+ & \sqrt{2}\Delta^{++} \\ v_\Delta + \delta + i\xi & -\Delta^+ \end{pmatrix}. \quad (11.48)$$

where  $\Delta^+$  is a singly-charged field,  $\Delta^{++}$  is a doubly-charged field,  $\delta$  is a neutral CP-even scalar,  $\xi$  is a neutral CP-odd scalar, and  $v_\Delta$  is the triplet VEV. The general scalar potential mixes the doublet and triplet components. After electroweak symmetry breaking there are seven physical mass eigenstates, denoted  $H^{\pm\pm}, H^\pm, A, H$ , and  $h$ .

A distinguishing feature of the HTM is that it violates the custodial symmetry of the SM; thus the  $\rho$  parameter deviates from 1 even at tree level. Letting  $x$  denote the ratio of triplet and doublet VEVs, the tree level expression [363] is:

$$\rho = \frac{1 + 2x^2}{1 + 4x^2}. \quad (11.49)$$

The measured value of the  $\rho$  parameter then limits [364] the triplet VEV to be quite small,  $x \lesssim 0.03$ , or  $v_\Delta < 8$  GeV. This constraint severely limits the role of the triplet scalar in the EWSB mechanism.

The small VEV of the Higgs triplet in the HTM is a virtue from the point of view of generating neutrino masses without the necessity for introducing right-handed neutrino fields. The gauge invariant dimension four interaction

$$h_{\nu ij} \ell_i^T C^{-1} i\sigma_2 \Delta \ell_j, \quad (11.50)$$

where  $\ell_i$  are the lepton doublets,  $C$  is the charge conjugation matrix, and  $h_{\nu ij}$  is a complex symmetric coupling matrix, generates a Majorana mass matrix for the neutrinos:

$$m_{\nu ij} = \sqrt{2} h_{\nu ij} v_\Delta. \quad (11.51)$$

This can be combined with the usual neutrino seesaw to produce what is known as the type-II seesaw [365].

The HTM suggests the exciting possibility of measuring parameters of the neutrino mass matrix at the LHC. If the doubly-charged Higgs is light enough and/or its couplings to  $W^+W^+$  are sufficiently suppressed, then its primary decay is into same-sign lepton pairs:  $H^{++} \rightarrow \ell_i^+ \ell_j^+$ ; from Eq. (11.50) and Eq. (11.51) it is apparent that these decays are in general lepton-flavor violating with branchings proportional to elements of the neutrino mass matrix [366].

Precision electroweak data constrain the mass spectrum as well as the triplet VEV of the HTM [363, 367, 368]. As described in Ref. [368], these constraints favor a spectrum where  $H^{++}$  is the lightest of the exotic bosons, and where the mass difference between  $H^+$  and  $H^{++}$  is a few hundred GeV. The favored triplet VEV is a few GeV, which also favors  $H^{++}$  decays into  $W^+W^+$  over same-sign dileptons.

The Georgi–Machacek model addresses the  $\rho$  parameter constraint directly by building in custodial symmetry. Writing the complex scalar doublet of the SM as a  $(2, 2)$  under  $SU(2)_L \times SU(2)_R$ , it is obvious that the next simplest construction respecting custodial symmetry is a scalar transforming like a  $(3, 3)$  [369]. These nine real degrees of freedom correspond to a complex electroweak triplet combined with a real triplet, with the scalar potential required to be invariant under

$SU(2)_R$ . Under the custodial  $SU(2)_{L+R}$ , they transform as  $1 \oplus 3 \oplus 5$ , with a CP-even neutral scalar as the custodial singlet (thus matching the SM Higgs boson), a CP-odd neutral scalar in the custodial triplet, and another CP-even neutral scalar in the custodial 5-plet.

The scalar components can be decomposed as [370]

$$\Xi = \begin{pmatrix} \chi_3^* & \xi_1 & \chi_1 \\ -\chi_2^* & \xi_2 & \chi_2 \\ \chi_1^* & -\xi_1^* & \chi_3 \end{pmatrix}, \quad (11.52)$$

where  $\xi_2$  is a real scalar and the others are complex scalars. Linear combinations of these account for the neutral custodial singlet, a neutral and singly-charged field making up the custodial triplet, and neutral, singly-charged, and doubly-charged fields making up the custodial 5-plet.

When combined with the usual SM doublet field  $\Phi$ , the electroweak scale  $v$  is now related to the doublet and triplet VEVs by

$$v^2 = v_\Phi^2 + 8v_\Xi^2. \quad (11.53)$$

Note that the GM triplets by themselves are sufficient to explain electroweak symmetry breaking and the existence of a 125 GeV neutral boson along with a custodial triplet of Goldstone bosons; the complex doublet field in the GM model is required to generate fermion masses via the usual dimension four Yukawa couplings. This raises the question of whether one can rule out the possibility that the 125 GeV boson is the neutral member of a custodial 5-plet rather than a custodial singlet, without invoking decays to fermions. A conclusive answer is given by observing that the ratio of the branching fractions to  $W$  versus  $Z$  bosons is completely determined by the custodial symmetry properties of the boson. For a custodial 5-plet, the ratio of the signal strength to  $WW$  over that to  $ZZ$  is predicted to be 1/4 that of a SM Higgs boson [369, 371], and thus already ruled out by the experimental results presented in Section VI.

Another interesting general feature of Higgs triplet models is that, after mixing, the SM-like neutral boson can have stronger couplings to  $WW$  and  $ZZ$  than predicted by the SM [362, 372]; this is in contrast to mixing with additional doublets and singlet, which can only reduce the  $WW$  and  $ZZ$  couplings versus the SM. This emphasizes that LHC Higgs data cannot extract model independent coupling strengths for the Higgs boson [230, 373].

Because of the built-in custodial symmetry, the triplet VEV in the GM model can be large compared to the doublet VEV. The custodial singlet neutral boson from the triplets mixes with the neutral boson from the doublet. Two interesting special cases are (i) the triplet VEV is small and the 125 GeV boson is SM-like except for small deviations, and (ii) the 125 GeV boson is mostly the custodial singlet neutral boson from the electroweak triplets. The phenomenology of the doubly-charged and singly-charged bosons is similar to that of the HTM. The constraints on the GM model from precision electroweak data, LEP data, and current LHC data are described in Refs. [370, 374–377].

## VII.6. Composite Higgs models

Within the SM, EWSB is posited but has no dynamical origin. Furthermore, the Higgs boson appears to be unnaturally light. A scenario that remedies these two catches is to consider the Higgs boson as a bound state of new dynamics becoming strong around the weak scale. The Higgs boson can be made significantly lighter than the other resonances of the strong sector if it appears as a pseudo-Nambu–Goldstone boson, see Refs. [15, 16, 378] for recent reviews.

### VII.6.1. Little Higgs models

The idea behind the Little Higgs models [379, 380] is to identify the Higgs doublet as a (pseudo) Nambu–Goldstone boson while keeping some sizable non-derivative interactions, in particular a largish Higgs quartic interaction. By analogy with QCD where the pions  $\pi^{\pm,0}$  appear as Nambu–Goldstone bosons associated to the breaking of the chiral symmetry  $SU(2)_L \times SU(2)_R/SU(2)$ , switching on some

interactions that break explicitly the global symmetry will generate masses for the would-be massless Nambu–Goldstone bosons of the order of  $g\Lambda_{G/H}/(4\pi)$ , where  $g$  is the coupling of the symmetry breaking interaction and  $\Lambda_{G/H} = 4\pi f_{G/H}$  is the dynamical scale of the global symmetry breaking  $G/H$ . In the case of the Higgs boson, the top Yukawa interaction or the gauge interactions themselves will certainly break explicitly (part of) the global symmetry since they act non-linearly on the Higgs boson. Therefore, obtaining a Higgs mass around 100 GeV would demand a dynamical scale  $\Lambda_{G/H}$  of the order of 1 TeV, which is known to lead to too large oblique corrections. Raising the strong dynamical scale by at least one order of magnitude requires an additional selection rule to ensure that a Higgs mass is generated at the 2-loop level only

$$m_H^2 = \frac{g^2}{16\pi^2} \Lambda_{G/H}^2 \rightarrow m_H^2 = \frac{g_1^2 g_2^2}{(16\pi^2)^2} \Lambda_{G/H}^2 \quad (11.54)$$

The way to enforce this selection rule is through a “collective breaking” of the global symmetry:

$$\mathcal{L} = \mathcal{L}_{G/H} + g_1 \mathcal{L}_1 + g_2 \mathcal{L}_2. \quad (11.55)$$

Each interaction  $\mathcal{L}_1$  or  $\mathcal{L}_2$  individually preserves a subset of the global symmetry such that the Higgs remains an exact Nambu–Goldstone boson whenever either  $g_1$  or  $g_2$  is vanishing. A mass term for the Higgs boson can be generated only by diagrams involving simultaneously both interactions. At one-loop, such diagram are not quadratically divergent, so the Higgs mass is not UV sensitive. Explicitly, the cancellation of the SM quadratic divergences is achieved by a set of new particles around the Fermi scale: gauge bosons, vector-like quarks, and extra massive scalars, which are related, by the original global symmetry, to the SM particles with the same spin. Contrary to supersymmetry, the cancellation of the quadratic divergences is achieved by same-spin particles. These new particles, with definite couplings to SM particles as dictated by the global symmetries of the theory, are perfect goals for the LHC.

The simplest incarnation of the collective breaking idea, the so-called littlest Higgs model, is based on a non-linear  $\sigma$ -model describing the spontaneous breaking  $SU(5)$  down to  $SO(5)$ . A subgroup  $SU(2)_1 \times U(1)_1 \times SU(2)_2 \times U(1)_2$  is weakly gauged. This model contains a weak doublet, that is identified with the Higgs doublet, and a complex weak triplet whose mass is not protected by collective breaking. Other popular little Higgs models are based on different coset spaces: minimal moose ( $SU(3)^2/SU(3)$ ) [381], the simplest little Higgs ( $SU(3)^2/SU(2)^2$ ) [382], the bestest little Higgs ( $SO(6)^2/SO(6)$ ) [383] *etc.* For comprehensive reviews, see Refs. [384, 385].

Generically, oblique corrections in Little Higgs models are reduced either by increasing the coupling of one of the gauge groups (in the case of product group models) or by increasing the masses of the  $W$  and  $Z$  partners, leading ultimately to a fine-tuning of the order of a few percents (see for instance Ref. [386] and references therein). The compatibility of Little Higgs models with experimental data is significantly improved when the global symmetry involves a custodial symmetry as well as a  $T$ -parity [387] under which, in analogy with  $R$ -parity in SUSY models, the SM particles are even and their partners are odd. Such Little Higgs models would therefore appear in colliders as jet(s) with missing transverse energy [388] and the ATLAS and CMS searches for squarks and gluinos [389] can be recast to obtain limits on the masses of the heavy vector-like quarks. The  $T$ -even top partner, with an expected mass below 1 TeV to cancel the top loop quadratic divergence without too much fine-tuning, would decay dominantly into a  $t + Z$  pair or into a  $b + W$  pair or even into  $t + H$ . The latest CMS and ATLAS direct searches [390] for vector-like top partners put a lower bound around 700 GeV on their mass, excluding the most natural region of the parameter space of these models, i.e., there is still fine-tuning at the per cent level.

The motivation for Little Higgs models is to solve the little hierarchy problem, i.e., to push the need for new physics (responsible for the stability of the weak scale) up to around 10 TeV. Per se, Little Higgs models are effective theories valid up to their cutoff scale  $\Lambda_{G/H}$ . Their UV completions could either be weakly or strongly coupled.



**Table 11.18:** Global symmetry breaking patterns and the corresponding Goldstone boson contents of the SM, the minimal composite Higgs model, the next to minimal composite Higgs model, and the minimal composite two Higgs doublet model. Note that the SU(3) model does not have a custodial invariance.  $a$  denotes a CP-odd scalar while  $h$  and  $H$  are CP-even scalars.

Model	Symmetry Pattern	Goldstone's
SM	SO(4)/SO(3)	$W_L, Z_L$
–	SU(3)/SU(2)×U(1)	$W_L, Z_L, H$
MCHM	SO(5)/SO(4)	$W_L, Z_L, H$
NMCHM	SO(6)/SO(5)	$W_L, Z_L, H, a$
MC2HM	SO(6)/SO(4)×SO(2)	$W_L, Z_L, h, H, H^\pm, a$

### VII.6.2. Models of partial compositeness

The Higgs boson is a special object. Even in composite models, it cannot appear as a regular resonance of the strong sector without endangering the viability of the setup when confronted to data. The way out is that the Higgs appears as a pseudo Nambu–Goldstone boson: the new strongly coupled sector is supposed to be invariant under a global symmetry  $G$  spontaneously broken to a subgroup  $H$  at the scale  $f$ . To avoid conflict with EW precision measurements, it is better if the strong interactions themselves do not break the EW symmetry, hence the SM gauge symmetry itself should be contained in  $H$ . See Table 11.18 for a few examples of coset spaces.

The SM (light) fermions and gauge bosons cannot be part of the strong sector itself since LEP data have already put stringent bounds on the compositeness scale of these particles far above the TeV scale. The gauge bosons couple to the strong sector by a weak gauging of an SU(2)×U(1) subgroup of the global symmetry  $G$ . Inspiration for the construction of such models comes from the AdS/CFT correspondence: the components of a gauge field along extra warped space dimension can be interpreted as the Goldstone boson resulting from the breaking of global symmetry of the strong sector. The couplings of the SM fermions to the strong sector could a priori take two different forms: (i) a bilinear coupling of two SM fermions to a composite scalar operator,  $\mathcal{O}$ , of the form  $\mathcal{L} = y \bar{q}_L u_R \mathcal{O} + \text{hc}$  in simple analogy with the SM Yukawa interactions. This is the way fermion masses were introduced in Technicolor theories and it generically comes with severe flavor problems and calls for extended model building gymnastics [391] to circumvent them; (ii) a linear mass mixing with fermionic vector-like operators:  $\mathcal{L} = \lambda_L \bar{q}_L \mathcal{Q}_R + \lambda_R \bar{U}_L u_R$ .  $\mathcal{Q}$  and  $\mathcal{U}$  are two fermionic composite operators of mass  $M_Q$  and  $M_U$ . Being part of the composite sector, they can have a direct coupling of generic order  $Y_*$  to the Higgs boson. In analogy with the photon- $\rho$  mixing in QCD, once the linear mixings are diagonalized, the physical states are a linear combination of elementary and composite fields. Effective Yukawa couplings are generated and read for instance for the up-type quark

$$y = Y_* \sin \theta_L \sin \theta_R \quad (11.56)$$

where  $\sin \theta_i = \lambda_i / \sqrt{M_U^2 + \lambda_i^2}$ ,  $i = L, R$ , measure the amount of compositeness of the SM left- and right-handed up-type quark. If the strong sector is flavor-anarchic, i.e., if the couplings of the Higgs to the composite fermions does not exhibit any particular flavor structure, the relation Eq. (11.56) implies that the light fermions are mostly elementary states ( $\sin \theta_i \ll 1$ ), while the third generation quarks need to have a sizable degree of compositeness. The partial compositeness paradigm offers an appealing dynamical explanation of the hierarchies in the fermion masses. In fact, assuming the strong sector to be almost conformal above the confinement scale, the low-energy values of the mass-mixing parameters  $\lambda_{L,R}$  are determined by the (constant) anomalous dimension of the composite operator they mix with. If the UV scale at which the linear mixings are generated is large, then  $\mathcal{O}(1)$  differences in the anomalous dimensions can generate naturally large hierarchies in the fermion masses via renormalization group running [392]. While the introduction of partial compositeness greatly ameliorated the flavor problem of the original composite Higgs models, nevertheless it did not solve the issue completely, at least in the case where the strong sector is assumed to be flavor-anarchic [393]. While

the partial compositeness set-up naturally emerges in models built in space-times with extra dimensions, no fully realistic microscopic realization of partial compositeness has been proposed in the literature.

Another nice aspect of the partial compositeness structure is the dynamical generation of the Higgs potential. The Higgs being a pseudo-Nambu–Goldstone boson, its mass does not receive any contribution from the strong sector itself but it is generated at the one-loop level via the couplings of the SM particles to the strong sector since these interactions are breaking the global symmetries under which the Higgs doublet transforms non-linearly. The leading contribution to the potential arises from top loops and it takes the form

$$V(H) = m_\rho^4 \frac{\sin \theta_{tL} \sin \theta_{tR}}{16\pi^2} (\alpha \cos(H/f) + \beta \sin^2(H/f) + \gamma \sin^4(H/f)), \quad (11.57)$$

where  $\alpha, \beta, \gamma$  are numbers of order 1 subject to selection rules following the transformation properties of the top quark under the global symmetries of the strong sector<sup>10</sup>, and  $m_\rho \approx g_\rho f$  is the typical mass scale of the strong sector resonances. The gauge contribution to the potential takes the form ( $g$  denotes the SU(2) gauge coupling)

$$m_\rho^4 \frac{g^2/g_\rho^2}{16\pi^2} \sin^2(H/f), \quad (11.58)$$

which is parametrically suppressed with respect to the top contribution by  $g^2/(g_\rho y_t)$ . The gauge term is always positive, and cannot trigger EWSB by itself. When  $\alpha = 0$ , the minimization condition of the potential simply reads

$$\sin^2 \frac{H}{f} = -\frac{\beta}{2\gamma}, \quad (11.59)$$

which implies that the natural expectation is that the scale  $f$  is generically of the order of the weak scale. Obtaining  $v \ll f$ , as required phenomenologically, requires some degree of tuning, which scales like  $\xi \equiv v^2/f^2$ . A mild tuning of the order of 10% ( $\xi \approx 0.1$ ) is typically enough to comply with electroweak precision constraints. This is an important point: in partial compositeness models, the entire Higgs potential is generated at one loop, therefore the separation between  $v$  and  $f$  can only be obtained at a price of a tuning. This marks a difference with respect to the Little Higgs models, which realize a parametric hierarchy between the quartic and mass terms through the collective symmetry breaking mechanism. In fact in Little Higgs models, the quartic coupling is a tree-level effect, leading to a potential

$$V(H) \approx \frac{g_{\text{SM}}^2}{16\pi^2} m_\rho^2 H^2 + g_{\text{SM}}^2 H^4, \quad (11.60)$$

where  $g_{\text{SM}}$  generically denotes the SM couplings. The minimization condition now reads  $v^2/f^2 \sim g_\rho^2/(16\pi^2)$ , therefore  $v$  is formally loop suppressed with respect to  $f$ . This is the major achievement of the Little Higgs constructions, which however comes at the price of the presence of sub-TeV vectors carrying EW quantum numbers and therefore giving rise generically to large oblique corrections to the propagators of the  $W$  and the  $Z$  gauge bosons.

After minimization, the potential Eq. (11.57) leads to an estimate of the Higgs mass as

$$m_H^2 \approx g_\rho^3 y_t^2 \pi^2 v^2. \quad (11.61)$$

It follows that the limit  $f \rightarrow \infty$ , i.e.  $\xi \rightarrow 0$ , is a true decoupling limit: all the resonances of the strong sector become heavy but the Higgs whose mass is protected by the symmetries of the coset  $G/H$ . When compared to the experimentally measured Higgs mass, this estimate puts an upper bound on the strength of the strong

<sup>10</sup> For instance in the SO(5)/SO(4) composite models, when the top quark is embedded into a spinorial representation of SO(5), then  $\gamma = 0$  and when it is part of a **5**, **10** or **14** representation,  $\alpha = 0$  as it can be inferred by looking at the structure of the  $H$ -dependent invariants built out of these representations [394]. The coefficient  $\gamma$  also generically comes with an extra power of the top compositeness fractions.



interactions:  $g_\rho \lesssim 2$ . In this limit of not so large coupling, the Higgs potential receives additional contributions. In particular, the fermionic resonances in the top sector which follow from the global symmetry structure of the new physics sector can help raising the Higgs mass. For instance in the minimal  $\text{SO}(5)/\text{SO}(4)$  model, using some dispersion relation techniques, one obtains [395]

$$m_H^2 \approx \frac{6}{\pi^2} \frac{m_t^2}{f^2} \frac{m_{Q_4}^2 m_{Q_1}^2}{m_{Q_1}^2 - m_{Q_4}^2} \log \left( \frac{m_{Q_1}}{m_{Q_4}} \right) \quad (11.62)$$

where  $Q_4$  and  $Q_1$  are fermionic color resonances transforming as a weak bi-doublet of hypercharge  $Y = 1/6$  and  $Y = 7/6$  and a weak singlet with hypercharge  $Y = -1/3$ . Therefore a 125 GeV mass can be obtained if at least one of the fermionic resonances is lighter than  $\sim 1.4 f$ . As in supersymmetric scenarios, the top sector is playing a crucial role in the dynamics of EWSB and can provide the first direct signs of new physics. The direct searches for these top partners, in particular the ones with exotic electric charges  $5/3$ , are already exploring the natural parameter spaces of these models [390, 396, 397].

The main physics properties of a pseudo Nambu–Goldstone Higgs boson can be captured in a model-independent way by a few number of higher-dimensional operators. Indeed, the strong dynamics at the origin of the composite Higgs singles out a few operators among the complete list presented earlier in Section VI: these are the operators that involve extra powers of the Higgs doublets and they are therefore generically suppressed by a factor  $1/f^2$  as opposed to the operators that involve extra derivatives or gauge bosons and are suppressed by a factor  $1/(g_\rho^2 f^2)$ . The relevant effective Lagrangian describing a strongly interacting light Higgs is:

$$\begin{aligned} \mathcal{L}_{\text{SILH}} = & \frac{c_H}{2f^2} \left( \partial_\mu (\Phi^\dagger \Phi) \right)^2 + \frac{c_T}{2f^2} \left( \Phi^\dagger \overleftrightarrow{D}^\mu \Phi \right)^2 - \frac{c_6 \lambda}{f^2} (\Phi^\dagger \Phi)^3 \\ & + \left( \sum_f \frac{c_f y_f}{f^2} \Phi^\dagger \Phi \bar{f}_L \Phi f_R + \text{h.c.} \right). \end{aligned} \quad (11.63)$$

Typically, these new interactions induce deviations in the Higgs couplings that scale like  $\mathcal{O}(v^2/f^2)$ , hence the measurements of the Higgs couplings can be translated into some constraints on the compositeness scale,  $4\pi f$ , of the Higgs boson. The peculiarity of these composite models is that, due to the Goldstone nature of the Higgs boson, the direct couplings to photons and gluons are further suppressed and generically the coupling modifiers defined in Section VI scale like

$$\begin{aligned} \kappa_{W,Z,f} & \sim 1 + \mathcal{O} \left( \frac{v^2}{f^2} \right), \\ \kappa_{Z\gamma} & \sim \mathcal{O} \left( \frac{v^2}{f^2} \right), \\ \kappa_{\gamma,g} & \sim \mathcal{O} \left( \frac{v^2}{f^2} \times \frac{y_t^2}{g_\rho^2} \right), \end{aligned} \quad (11.64)$$

where  $g_\rho$  denotes the typical coupling strength among the states of the strongly coupled sector and  $y_t$  is the top Yukawa coupling, the largest interaction that breaks the Goldstone symmetry. The  $\kappa_{Z\gamma, \gamma, g}$  coupling modifiers are not generated by the strong coupling operators of Eq. (11.63) but some subleading form-factor operator generated by loops of heavy resonances of the strong sector. The coupling modifiers also receive additional contributions from the other resonances of the strong sector, in particular the fermionic resonances of the top sector that are required to be light to generate a 125 GeV Higgs mass. Some indirect information on the resonance spectrum could thus be inferred by a precise measurement of the Higgs coupling deviations. However, it was realized [398] that the task is actually complicated by the fact that, in the minimal models, these top partners give a contribution to both  $\kappa_t$  (resulting from a modification of the top Yukawa coupling) and  $\kappa_\gamma$  and  $\kappa_g$  (resulting from new heavy particles running into the loops) and the structure of interactions are such that the net effect vanishes for inclusive quantities like  $\sigma(gg \rightarrow H)$  or  $\Gamma(H \rightarrow \gamma\gamma)$  as a consequence of the Higgs low energy theorem [24, 25, 231]. So one would need to rely on differential distribution, like the Higgs  $p_T$

distribution [399], to see the top partner effects in Higgs data [400]. The off-shell channel  $gg \rightarrow h^* \rightarrow 4\ell$  [401] and the double Higgs production  $gg \rightarrow hh$  [402] can also help to resolve the gluon loop and separate the top and top-partner contributions.

### VII.6.3. Minimal composite Higgs models

The minimal composite Higgs models (MCHM) are concrete examples of the partial compositeness paradigm. The Higgs doublet is described by the coset space  $\text{SO}(5)/\text{SO}(4)$  where a subgroup  $\text{SU}(2)_L \times \text{U}(1)_Y$  is weakly gauged under which the four Goldstone bosons transform as a doublet of hypercharge 1. There is some freedom on how the global symmetry is acting on the SM fermions: in MCHM4 [394] the quarks and leptons are embedded into spinorial representations of  $\text{SO}(5)$ , while in MCHM5 [403] they are part of fundamental representations (it might also be interesting phenomenologically to consider larger representations like MCHM14 [404] with the SM fermions inside a representation of dimension 14). The non-linearly realized symmetry acting on the Goldstone bosons leads to general predictions of the coupling of the Higgs boson to the EW gauge bosons. For instance, it can be shown that the quadratic terms in the  $W$  and  $Z$  bosons read

$$m_W^2(H) \left( W_\mu W^\mu + \frac{1}{2 \cos^2 \theta_W} Z_\mu Z^\mu \right)$$

with  $m_W(H) = \frac{gf}{2} \sin \frac{H}{f}$ . Expanding around the EW vacuum, the expression of the weak scale is:

$$v = f \sin(\langle H \rangle / f), \quad (11.65)$$

and the values of the modified Higgs couplings to the  $W$  and  $Z$ :

$$g_{HVV} = \frac{2m_V^2}{v} \sqrt{1 - v^2/f^2}, \quad g_{HHVV} = \frac{2m_V^2}{v^2} (1 - 2v^2/f^2). \quad (11.66)$$

Note that the Higgs couplings to gauge bosons is always suppressed compared to the SM prediction. This is a general result [405] that holds as long as the coset space is compact.

The Higgs couplings to the fermions depend on the representation which the SM fermions are embedded into. For the most commonly used embeddings, they take the following forms

$$\begin{aligned} \text{MCHM4} : g_{Hff} &= \frac{m_f}{v} \sqrt{1 - v^2/f^2}, \\ \text{MCHM5} : g_{Hff} &= \frac{m_f}{v} \frac{1 - 2v^2/f^2}{\sqrt{1 - v^2/f^2}}, \\ \text{MCHM14} : g_{Hff} &= \frac{m_f}{v} \left( 1 + A(M_{1,4,9}) \frac{v^2}{f^2} + \mathcal{O}(v^4/f^4) \right), \\ &\text{with } A(M_{1,4,9}) = \frac{3M_1 M_4 - 11M_1 M_9 + 8M_4 M_9}{2M_9(M_1 - M_4)}. \end{aligned} \quad (11.67)$$

While, in MCHM4 and MCHM5, the modifications of the couplings depend only on the Higgs compositeness scale, in MCHM14 the leading corrections depend also on the mass spectrum of the resonances parametrized by  $M_1, M_4$  and  $M_9$  [404]. This is due to the fact that more than one  $\text{SO}(5)$  invariant gives rise to SM fermion masses. The  $(\kappa_V, \kappa_f)$  experimental fit of the Higgs couplings can be used to derive a lower bound on the Higgs compositeness scale  $4\pi f \gtrsim 9 \text{ TeV}$ , which is less stringent than the indirect bound obtained from EW precision data,  $4\pi f \gtrsim 15 \text{ TeV}$  [406] but more robust and less subject on assumptions [407].

### VII.6.4. Twin Higgs models

In all composite models presented above, the particles responsible for canceling the quadratic divergences in the Higgs mass are charged under the SM gauge symmetries. In particular, the top partner carries color charge, implying a reasonably large minimal production cross section at the LHC. An alternative scenario, which is experimentally quite challenging and might explain the null result in various new physics searches, is the case nowadays referred to

as “neutral naturalness” [20, 21], where the particles canceling the 1-loop quadratic divergences are neutral under the SM. The canonical example for such theories is the Twin Higgs model of [20]. This is an example of a pseudo-Goldstone boson Higgs theory, with an approximate global  $SU(4)$  symmetry broken to  $SU(3)$ . The Twin Higgs model is obtained by gauging the  $SU(2)_A \times SU(2)_B$  subgroup of  $SU(4)$ , where  $SU(2)_A$  is identified with the SM  $SU(2)_L$ , while  $SU(2)_B$  is the twin  $SU(2)$  group. Gauging this subgroup breaks the  $SU(4)$  symmetry explicitly, but quadratically divergent corrections given do not involve the Higgs boson when the gauge couplings of the two  $SU(2)$  subgroups are equal,  $g_A = g_B$ . The  $SU(4) \rightarrow SU(3)$  breaking will also result in the breaking of the twin  $SU(2)_B$  group and as a result three of the seven Goldstone bosons will be eaten, leaving 4 Goldstone bosons corresponding to the SM Higgs doublet  $h$ . In fact imposing the  $Z_2$  symmetry on the full model will ensure the cancellation of all 1-loop quadratic divergences to the Higgs mass. Logarithmically divergent terms can however arise for example from gauge loops, leading to a Higgs mass of order  $g^2 f / 4\pi$ , which is of the order of the physical Higgs mass for  $f \sim 1$  TeV. The quadratic divergences from the top sector can be eliminated if the  $Z_2$  protecting the Higgs mass remains unbroken by the couplings that result in the top Yukawa coupling. This can be achieved by introducing top partners charged under a twin  $SU(3)_c$ . In this case the quadratic divergences are cancelled by top partners that are neutral under the SM gauge symmetries.

Twin Higgs models are low-energy effective theories valid up to a cutoff scale of order  $\Lambda \sim 4\pi f \sim 5\text{--}10$  TeV, beyond which a UV completion has to be specified. The simplest such possibility is to also make the Higgs composite, and UV complete the twin-Higgs model via gauge and top partners at masses of the order of a few TeV. A concrete implementation is the holographic twin Higgs model [408], which also incorporates a custodial symmetry to protect the  $T$ -parameter from large corrections. It is based on a warped extra dimensional theory with a bulk  $SO(8)$  gauge group, which incorporates the  $SU(4)$  global symmetry discussed above enlarged to contain the  $SU(2)_L \times SU(2)_R$  custodial symmetry. In addition the bulk contains either a full  $SU(7)$  group or an  $SU(3) \times SU(3) \times U(1) \times U(1) \times Z_2$  subgroup of it to incorporate QCD, its twin, and hypercharge. The breaking on the UV brane is to the SM and the twin SM symmetries, while on the IR brane  $SO(8) \rightarrow SO(7)$ , giving rise to the 7 Goldstone bosons, three of which will be again eaten by the twin  $W, Z$ . The main difference compared to ordinary composite Higgs models is that in composite twin Higgs models the cancellation of the one-loop quadratic divergences is achieved by the twin partners of order 700 GeV–1 TeV, which are uncharged under the SM gauge group. This allows the IR scale of the warped extra dimension to be raised to the multi-TeV range without reintroducing the hierarchy problem. The role of the composite partners is to UV complete the theory, rather than the cancellation of the one-loop quadratic divergences. For more details about the composite twin Higgs models, see Refs. [409].

### VII.7. The Higgs boson as a dilaton

The possibility that the new particle  $H^0$  discovered at the LHC is in fact the Goldstone boson associated to the spontaneous breaking of scale invariance at a scale  $f$  attracted some attention [18, 19] but is now challenged by the fact that all its properties are in good agreement with those predicted for the SM Higgs. And this scenario now requires rather involved model-building engineering. The first issue is the fact that the observed scalar couplings are close to their SM values. In a generic theory of spontaneously broken scale invariance, order one shifts are possible, and indeed expected in most models. Also, the apparent hierarchy between the light scalar and the cutoff of the dilaton effective theory is not reconcilable with the general walking technicolor (or Higgsless) type scenario unless a tuning is imposed.

The general couplings of a wide class of dilaton models are given

(at leading order in a low-energy theorem limit for dilatons) by

$$\mathcal{L}_{\text{dilaton}} = \frac{\sigma}{f} \left[ 2M_W^2 W_\mu^\pm W^{\pm\mu} + M_Z^2 Z_\mu Z^\mu + \sum_{ij} \sqrt{m_f^i m_f^j} \Gamma^{ij} \bar{\psi}^i \psi^j \right] + \frac{\sigma}{f} \left[ \frac{2}{e} \Delta\beta^{\text{em}} F_{\mu\nu}^2 + \frac{2}{g_3} \Delta\beta^{\text{QCD}} G_{\mu\nu}^a{}^2 \right] \quad (11.68)$$

where  $\Gamma^{ij}$  is a matrix that depends upon anomalous dimensions of operators in the conformal theory that give rise to fermion masses, and the terms  $\Delta\beta$  are the differences in the beta functions of electromagnetism or QCD at scales above and below the scale at which conformal symmetry is spontaneously broken. The SM low energy theorem limit for the SM Higgs is obtained from this expression by taking

$$f = v, \quad \Gamma^{ij} = I_{3 \times 3}, \quad \Delta\beta^{\text{em}} = \beta_{\text{top}}^{\text{em}} + \beta_W^{\text{em}}, \quad \Delta\beta^{\text{QCD}} = \beta_{\text{top}}^{\text{QCD}}. \quad (11.69)$$

It is unclear why these relations should be approximately realized in a generic conformal field theory, as must be the case to be consistent with current data and allow for a scalar with mass of about 125 GeV. For example, in warped models of electroweak symmetry breaking (AdS/CFT duals to theories with spontaneously approximate broken conformal invariance), the ratio  $v/f$  is a function of the geometry, and is suppressed when the 5D theory is perturbative, contrary to the experimental result that the  $v/f$  ratio should be close to 1, implying that the underlying CFT may not be a large  $N$  CFT.

An additional complication is that the mass of the dilaton is expected to be, along with many other resonances, around the cutoff scale of the strongly interacting theory responsible for breaking the scale invariance spontaneously. Suppression of the dilation mass either requires a tuning of order  $v^2/\Lambda^2 \sim$  percent, or a very special conformal dynamics where the beta function of the interaction leading to the scale invariance breaking remains small over a large region of couplings [410].

### VII.8. Searches for signatures of extended Higgs sectors

The measurements described in Section III have established the existence of one state of the electroweak symmetry breaking sector, compatible with a SM Higgs boson, but not that it is the only one.

Various classes of models beyond the SM discussed above require extended Higgs sectors. These models, and in particular the MSSM and the NMSSM serve as guiding principle of the experimental searches for additional scalar states beyond the SM. However these searches are made as model-independent as possible and can be summarized in the following classes: (i) the search for an additional CP-even state mostly in the high mass domain decaying to vector bosons, which would correspond to the heavy CP-even state in a generic 2HDM where the light state would be the discovered  $H$  or a generic additional singlet; (ii) the search for a state in the high mass domain decaying to pairs of fermions, which would correspond a CP-odd  $A$  and the heavy CP-even state  $H$  in a generic 2HDM; (iii) the search for charged Higgs bosons, which also appear in generic 2HDMs; (iv) the search for a CP-odd state  $a$  in the low mass region which appears in the NMSSM; and (v) doubly charged Higgs which are motivated in extensions of the Higgs sector with triplets.

#### (i) Searches for an additional CP-even state

##### (a) Exclusion limits from LEP

The LEP searches for the SM Higgs boson put a lower limit of 114 GeV on its mass, but also have relevance for non-SM Higgs bosons. These searches were also interpreted as 95% CL upper bounds on the ratio of the coupling  $g_{HZZ}$  to its SM prediction as a function of the Higgs boson mass [125]. Among the MSSM new benchmarks, the low- $m_H$  is one example which is disfavored by these searches at low mass, and nearly ruled out by current direct constraints and charged Higgs limits from LHC. Another example is the light CP-even Higgs boson of the NMSSM which is constrained to project predominantly onto the EW singlet component. An additional motivation for these

**Table 11.19:** Summary of references to searches for additional states from extended Higgs sectors, where (BBr) denotes the BaBar experiment, (TeV) the Tevatron experiments.

	ATLAS	CMS	Other experiments
CP-even $H$			
$H \rightarrow \gamma\gamma$	—	[415]	—
$H \rightarrow Z\gamma$	[171]	[170]	—
$H \rightarrow ZZ \rightarrow 4\ell$	[470]	[471]	—
$H \rightarrow ZZ \rightarrow \ell\ell\nu\nu$	[472]	[473]	—
$H \rightarrow ZZ \rightarrow \ell\ell q\bar{q}$	[474, 475]	[476]	—
$H \rightarrow WW \rightarrow \ell\nu\ell\nu$	[477]	[471]	—
$H \rightarrow WW \rightarrow \ell\nu\ell\nu$ (2HDM)	[414]	[471]	—
$H \rightarrow WW \rightarrow \ell\nu q\bar{q}$	[478]	[479, 480]	—
$H \rightarrow hh \rightarrow b\bar{b}\tau\tau, b\bar{b}\gamma\gamma, 4b, \gamma\gamma WW^*$	[412]	[413, 411]	—
CP-odd $A$ (and/or CP-even $H$ )			
$H, A \rightarrow \tau^+\tau^-$	[429]	[430]	[427, 428]-TeV [481]-LHCb
$H, A \rightarrow \mu^+\mu^-$	[429]	—	—
$H, A \rightarrow t\bar{t}$	[435]	—	—
$H, A \rightarrow b\bar{b}$	—	[431]	[425, 426]-TeV
$A \rightarrow hZ \rightarrow b\bar{b}\ell\ell, \ell\ell\tau\tau, \nu\bar{\nu}b\bar{b}$	[438]	[411]	—
Charged $H^\pm$			
$H^\pm \rightarrow \tau^\pm\nu$	[447]	[449]	—
$H^\pm \rightarrow cs$	[451]	[452]	—
$H^\pm \rightarrow tb$	[450]	[449]	—
$H^\pm \rightarrow W^\pm Z$	[453]	[449]	—
CP-odd NMSSM $a$			
$a \rightarrow \mu^+\mu^-$	[464]	[465]	—
$h \rightarrow aa \rightarrow 4\mu, 4\tau, 2\mu 2\tau, 4\gamma$	[482, 466]	[467, 468]	[456]-TeV, [469]-LEP
$\Upsilon_{1s,3s} \rightarrow a\gamma$	—	—	[461, 462]-BBr
Doubly Charged $H^\pm$	[483]	[484]	—

scenarios is given by the slight excess observed at LEP [125] at a Higgs boson mass hypothesis of approximately 98 GeV. The light CP-even Higgs boson  $h$  was also searched for in association with the CP-odd  $A$ , these searches are described in Section III.

#### (b) Searches at the LHC

The searches for the SM Higgs boson before the discovery covered a wide range of mass hypotheses. Until recently the range of investigation at LHC was from 100 GeV to 600 GeV. It has been extended to masses of up to 1 TeV. At the Tevatron this mass range was limited to up to 200 GeV. Since the discovery, the SM Higgs boson searches are reappraised to search for a heavy CP-even state. This state could be the heavy CP-even Higgs boson of a 2HDM, or a generic additional singlet. In both cases the natural width of the additional  $H$  state can be very different from that of the SM Higgs boson. To preserve unitarity of the longitudinal vector boson scattering and the longitudinal vector boson scattering into fermion pairs, the couplings of the additional CP-even Higgs boson to gauge bosons and fermions should not be too large and should constrain the natural width to be smaller than that of a unique Higgs boson at high mass with couplings to fermions and gauge bosons as predicted by the SM (and provided that trilinear and quartic couplings are not too large and that no new state affects the heavy state total width). It is therefore reasonable to consider total widths for the high mass CP-even state smaller than the equivalent SM width. For the sake of generality these searches should be done as a function of Higgs boson mass and total width. Until recently only two cases have been investigated: (i) the SM width using the complex pole scheme (CPS), and (ii) the narrow width approximation.

Searches for the Higgs boson in the  $H \rightarrow \gamma\gamma$ ,  $H \rightarrow Z\gamma$ ,  $H \rightarrow W^{(*)}W^{(*)}$  in the  $\ell\nu\ell\nu$  and  $\ell\nu q\bar{q}$  channels, and the  $H \rightarrow Z^{(*)}Z^{(*)}$

searches in the  $4\ell$ ,  $\ell\ell q\bar{q}$  and  $\ell\ell\nu\nu$  channels have also been done, but in most cases are simple reinterpretations of the SM Higgs search in the CPS scheme. Recent references are summarized in Table 11.19.

#### (d) Searches for an additional resonance decaying to a pair of $h$

In addition to the rare and expected Higgs pair production mode, high mass CP-even Higgs bosons can be searched for in the resonant double Higgs mode. Searches for such processes, where the Higgs boson is used as a tool for searches for new phenomena beyond the SM, have been carried out in four distinct modes depending on the subsequent decays of each Higgs boson. The ATLAS and CMS Collaborations have searched for the  $H \rightarrow hh \rightarrow b\bar{b}\tau\tau$  [411, 412] and  $b\bar{b}\gamma\gamma$  [413, 412] final states. The ATLAS Collaboration has also searched for the  $H \rightarrow hh \rightarrow 4b$  and  $\gamma\gamma WW^*$  final states [412]. For masses hypotheses of an additional Higgs boson below 500 GeV, the two dominant search channels are the  $b\bar{b}\gamma\gamma$  and the  $b\bar{b}\tau\tau$ . For masses above 500 GeV, the most powerful search channel is the  $4b$  final state. The ATLAS Collaboration has also performed a combination of these search results assuming that the Higgs boson has standard decay rates [412].

#### (d) Searches for an additional state with the presence of $h$

In the post-discovery era, analyses in general need to take into account the presence of the newly discovered state. For searches with sufficiently high resolution of additional states non degenerate in mass, the strength of the observed state and limits on the signal strength of a potential additional state can be set independently, as discussed in the next section. However in some cases, such as when a channel does not have a sufficiently fine mass resolution or when the states are nearly degenerate in mass, specific analyses need to be designed. There are two examples of such analyses: (i) the search for an additional state in the  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  channel in ATLAS and (ii) the search for nearly degenerate states in the  $H \rightarrow \gamma\gamma$  channel with the CMS detector.

The search in the  $H \rightarrow W^{(*)}W^{(*)} \rightarrow \ell\nu\ell\nu$  channel, for an additional state is done using a boosted decision tree combining several discriminating kinematic characteristics to separate the signal from the background and a high mass signal  $H$  from the lower mass state  $h$  [414]. A simultaneous fit of the two states  $h$  and  $H$  is then made to test the presence of an additional state. In this case, the usual null hypothesis of background includes including the SM signal.

The CMS search for nearly degenerate mass states decaying to a pair of photons [415] is more generic and could for instance apply to CP-odd Higgs bosons as well. It consists of a fit to the diphoton mass spectrum using two nearly degenerate mass templates.

#### (e) Type I 2HDM and fermiophobia

The measurements of coupling properties of  $H$  indirectly exclude that the discovered state is fermiophobic. However, the presence of an additional fermiophobic state, as predicted by Type I 2HDMs, is not excluded. Prior to the discovery, ATLAS and CMS have performed searches for a fermiophobic Higgs boson, *i.e.* produced through couplings with vector bosons only (VBF and VH) and decaying in  $h_f \rightarrow \gamma\gamma$ , optimized for fermiophobic signatures in the diphoton channel [416, 417]. CMS has further combined these results with searches for  $h_f \rightarrow W^+W^-$  and  $h_f \rightarrow ZZ$  assuming fermiophobic production and decay [418]. CMS excludes a fermiophobic Higgs boson in the range  $110 \text{ GeV} < m_H < 188 \text{ GeV}$  at the 95% C.L.

#### (f) Interpretation benchmarks in the light of the discovered Higgs boson

Two specific benchmark scenarios driven by unitarity relations are proposed in Ref. [44], assuming the existence of an additional state  $h'$  with coupling scale factors, *i.e.*, deviations from the couplings predicted for the SM Higgs at the same mass, denoted  $\kappa_V'$  and  $\kappa_F'$  for the couplings of  $h'$  to vector bosons and fermions respectively. The gauge boson scattering unitarity then yields the following sum rule

$$\kappa_V'^2 + \kappa_F'^2 = 1 \quad (11.70)$$

and the unitarization of the gauge boson scattering to fermions yields

$$\kappa_V \cdot \kappa_F + \kappa'_V \cdot \kappa'_F = 1 \quad (11.71)$$

The two benchmark scenarios are then defined as follows: (i) a single coupling scale factor is assumed for the gauge bosons and the fermions, with an additional parameter to take into account decays to new states; (ii) two parameters are used to describe independently the couplings to fermions and the couplings to vector bosons. A direct application of the latter can be done in the CP-even sector of the type-I 2HDM.

**(ii) Searches for additional neutral states ( $\phi \equiv h, H, A$ ) decaying to fermions**

*(a) Exclusion limits from LEP*

In  $e^+e^-$  collisions at LEP centre-of-mass energies, the main production mechanisms of the neutral MSSM Higgs bosons were the Higgs-strahlung processes  $e^+e^- \rightarrow hZ, HZ$  and the pair production processes  $e^+e^- \rightarrow hA, HA$ , while the vector boson fusion processes played a marginal role. Higgs boson decays to  $b\bar{b}$  and  $\tau^+\tau^-$  were used in these searches.

The searches and limits from the four LEP experiments are described in Refs. [419,420]. The combined LEP data did not contain any excess of events which would imply the production of a Higgs boson, and combined limits were derived [421]. For  $m_A \gg M_Z$  the limit on  $m_h$  is nearly that of the SM searches, as  $\sin^2(\beta - \alpha) \approx 1$ . For high values of  $\tan\beta$  and low  $m_A$  ( $m_A \leq m_h^{\max}$ ), the  $e^+e^- \rightarrow hA$  searches become the most important, and the lightest Higgs  $h$  is non SM-like. In this region, the 95% CL mass bounds are  $m_h > 92.8$  GeV and  $m_A > 93.4$  GeV. In the  $m_h$ -max. scenario, values of  $\tan\beta$  from 0.7 to 2.0 are excluded taking  $m_t = 174.3$  GeV, while a much larger  $\tan\beta$  region is excluded for other benchmark scenarios such as the no-mixing one.

A flavor-independent limit for Higgs bosons in the Higgs-strahlung process at LEP has also been set at 112 GeV [422].

Neutral Higgs bosons may also be produced by Yukawa processes  $e^+e^- \rightarrow f\bar{f}\phi$ , where the Higgs particle  $\phi \equiv h, H, A$ , is radiated off a massive fermion ( $f \equiv b$  or  $\tau^\pm$ ). These processes can be dominant at low masses, and whenever the  $e^+e^- \rightarrow hZ$  and  $hA$  processes are suppressed. The corresponding ratios of the  $f\bar{f}h$  and  $f\bar{f}A$  couplings to the SM coupling are  $\sin\alpha/\cos\beta$  and  $\tan\beta$ , respectively. The LEP data have been used to search for  $b\bar{b}b\bar{b}$ ,  $b\bar{b}\tau^+\tau^-$ , and  $\tau^+\tau^-\tau^+\tau^-$  final states [423,424]. Regions of low mass and high enhancement factors are excluded by these searches.

The searches for the Higgs boson at LEP also included the case where it does not predominantly decay to a pair of  $b$  quarks. All four collaborations conducted dedicated searches for the Higgs boson with reduced model dependence, assuming it is produced via the Higgs-strahlung process, and not addressing its flavor of decay, a lower limit on the Higgs mass of 112.9 GeV is set by combining the data of all four experiments [422].

Using an effective Lagrangian approach and combining results sensitive to the  $h\gamma\gamma$ ,  $hZ\gamma$  and  $hZZ$  couplings, an interpretation of several searches for the Higgs boson was made and set a lower limit of 106.7 GeV on the mass of a Higgs boson that can couple anomalously to photons [422].

*(b) Searches at the Tevatron and LHC*

The best sensitivity is in the regime with low to moderate  $m_A$  and with large  $\tan\beta$  which enhances the couplings of the Higgs bosons to down-type fermions. The corresponding limits on the Higgs boson production cross section times the branching ratio of the Higgs boson into down-type fermions can be interpreted in MSSM benchmark scenarios [259]. If  $\phi = A, H$  for  $m_A > m_h^{\max}$ , and  $\phi = A, h$  for  $m_A < m_h^{\max}$ , the most promising channels at the Tevatron are the inclusive  $p\bar{p} \rightarrow \phi \rightarrow \tau^+\tau^-$  process, with contributions from both  $gg \rightarrow \phi$  and  $b\bar{b}\phi$  production, and  $b\bar{b}\phi, \phi \rightarrow \tau^+\tau^-$  or  $\phi \rightarrow b\bar{b}$ , with  $b\tau\tau$  or three tagged  $b$ -jets in the final state, respectively. Although Higgs

boson production via gluon fusion has a higher cross section in general than via associated production, it cannot be used to study the  $\phi \rightarrow b\bar{b}$  decay mode since the signal is overwhelmed by the QCD background.

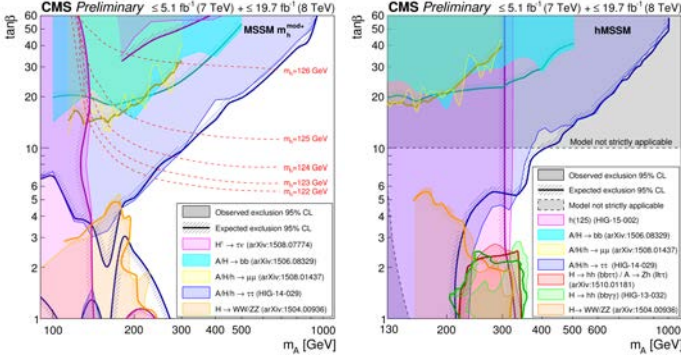
The CDF and D0 collaborations have searched for neutral Higgs bosons produced in association with bottom quarks and which decay into  $b\bar{b}$  [425,426], or into  $\tau^+\tau^-$  [427,428]. The most recent searches in the  $b\bar{b}\phi$  channel with  $\phi \rightarrow b\bar{b}$  analyze approximately  $2.6\text{ fb}^{-1}$  of data (CDF) and  $5.2\text{ fb}^{-1}$  (D0), seeking events with at least three  $b$ -tagged jets. The cross section is defined such that at least one  $b$  quark not from  $\phi$  decay is required to have  $p_T > 20$  GeV and  $|\eta| < 5$ . The invariant mass of the two leading jets as well as  $b$ -tagging variables are used to discriminate the signal from the backgrounds. The QCD background rates and shapes are inferred from data control samples, in particular, the sample with two  $b$ -tagged jets and a third, untagged jet. Separate-signal hypotheses are tested and limits are placed on  $\sigma(p\bar{p} \rightarrow b\bar{b}\phi) \times \text{BR}(\phi \rightarrow b\bar{b})$ . A local excess of approximately  $2.5\sigma$  significance has been observed in the mass range of 130–160 GeV, but D0's search is more sensitive and sets stronger limits. The D0 result had an  $\mathcal{O}(2\sigma)$  local upward fluctuation in the 110 to 125 GeV mass range. These results have been superseded by the LHC searches and the excess seen in the D0 experiment has not been confirmed elsewhere.

ATLAS and CMS also search for  $\phi \rightarrow \tau^+\tau^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV. ATLAS seeks tau pairs in  $4.7\text{--}4.8\text{ fb}^{-1}$  of data [429], and the search by CMS uses the full  $4.9\text{ fb}^{-1}$  of 7 TeV data  $4.9\text{ fb}^{-1}$  of 8 TeV data [430] and  $b\bar{b}$  [431]. The searches are performed in categories of the decays of the two tau leptons:  $e\tau_{\text{had}}, \mu\tau_{\text{had}}, e\mu$ , and  $\mu\mu$ , where  $\tau_{\text{had}}$  denotes a tau lepton which decays to one or more hadrons plus a tau neutrino,  $e$  denotes  $\tau \rightarrow e\nu\nu$ , and  $\mu$  denotes  $\tau \rightarrow \mu\nu\nu$ . The dominant background comes from  $Z \rightarrow \tau^+\tau^-$  decays, although  $t\bar{t}$ ,  $W$ +jets and  $Z$ +jets events contribute as well. Separating events into categories based on the number of  $b$ -tagged jets improves the sensitivity in the MSSM. The  $b\bar{b}$  annihilation process and radiation of a Higgs boson from a  $b$  quark gives rise to events in which the Higgs boson is accompanied by a  $b\bar{b}$  pair in the final state. Requiring the presence of one or more  $b$  jets reduces the background from  $Z$ +jets. Data control samples are used to constrain background rates. The rates for jets to be identified as a hadronically decaying tau lepton are measured in dijet samples, and  $W$ +jets samples provide a measurement of the rate of events that, with a fake hadronic tau, can pass the signal selection requirements. Lepton fake rates are measured using samples of isolated lepton candidates and same-sign lepton candidates. Constraints from the CMS searches for  $h \rightarrow \tau^+\tau^-$  and  $h \rightarrow b\bar{b}$  are shown in Fig. 11.25 in the  $m_h$ -mod+ scenario defined in [258] and in the hMSSM approximation defined in [307]. The neutral Higgs boson searches consider the contributions of both the CP-odd and CP-even neutral Higgs bosons with enhanced couplings to bottom quarks, similarly as it was done for the Tevatron results. In Fig. 11.25, decays of the charged Higgs into  $\tau\nu$  and of the heavy Higgs  $H$  decaying into a pair of SM-like Higgs bosons or gauge bosons, or of  $A$  decaying into  $hZ$  are also being constrained. In addition, decays of the neutral Higgs bosons into muon pairs are also being explored. Observe that in the  $m_h$ -mod+ scenario the region of  $\tan\beta$  lower than 5 does not allow for a Higgs mass  $m_h$  close to 125 GeV, as shown in the figure. For the hMSSM scenario, instead, the SM-like Higgs mass is fixed as an input and hence the requirement that it is close to 125 GeV is always fulfilled, although this may imply other limitations as discussed in section VII.1.1.

A search for  $\phi \rightarrow \mu^+\mu^-$  has also been performed by the ATLAS collaboration [429]. The exclusion limits obtained are given in terms of cross section times branching fraction and combined with those of  $\phi \rightarrow \tau^+\tau^-$  [429].

A search for pseudoscalar Higgs bosons at intermediate to low masses, below the  $Z$  mass (in the 25 GeV to 80 GeV mass range) has been performed by the CMS collaboration [433]. A light pseudoscalar in this mass range is excluded by current direct constraints in the MSSM but not in general 2HDMs [434]. This search is done in the decay channel where the pseudoscalar Higgs boson decays to a pair of taus and is produced in association with a pair of  $b$ -quarks.

Finally searches for a resonance decaying to a top quark pair were already done by ATLAS [435] and CMS [436]. These searches



**Figure 11.25:** The 95% CL exclusion contours in the  $(m_A, \tan \beta)$  parameter space for the hMSSM scenario (right panel) and for the  $m_h$ -mod+ scenario (left panel), for several search channels [432].

were interpreted as searches for scalar resonances by ATLAS [435], however an important missing component of these searches is an accurate treatment of the interference effects between the signal and the continuum background, that may have a non-conventional pure dip structure and can be non-negligible for a high mass state decaying to top quarks [357, 437].

The LHC has the potential to explore a broad range of SUSY parameter space through the search for non-SM-like Higgs bosons. Nevertheless, Fig. 11.25 shows a broad region with intermediate  $\tan \beta$  and large values of  $m_A$  that is not tested by present neutral or charged Higgs boson searches, and which cannot be covered completely via these searches, even with much larger data sets. In this region of parameter space it is possible that only the SM-like Higgs boson can be within the LHC's reach. If no other state of the EWSB sector than  $H$  is discovered, it may be challenging to determine only from the Higgs sector whether there is a supersymmetric extension of the SM in nature.

### (iii) Searches for a CP-odd state decaying to $hZ$

Similarly to the search for a CP-even high mass Higgs boson decaying to a pair of Higgs bosons, the search for a CP-odd states decaying  $hZ$  was carried out at the LHC by the ATLAS and CMS experiments. The ATLAS Collaboration has performed the search in three main final states corresponding to the following subsequent decays of the Higgs and Z bosons [438]:  $(Z \rightarrow \ell\ell)(h \rightarrow b\bar{b})$ ,  $(Z \rightarrow \nu\nu)(h \rightarrow b\bar{b})$  and  $(Z \rightarrow \ell\ell)(h \rightarrow \tau\tau)$ ; and their combination assuming SM Higgs decay branching fractions. The CMS Collaboration has performed a search in the  $(Z \rightarrow \ell\ell)(h \rightarrow \tau\tau)$  final state [411]. These searches have been used to constrain the parameter space of 2HDMs. In the MSSM these searches place limits on small values of  $\tan \beta$  for masses of  $A$  comprised between 220 GeV and 360 GeV as illustrated in Fig. 11.25.

### (iv) Searches for charged Higgs bosons $H^\pm$

At  $e^+e^-$  colliders charged Higgs bosons can be pair produced in the  $s$ -channel via  $\gamma$  or  $Z$  boson exchange. This process is dominant in the LEP centre-of-mass energies range *i.e.* up to 209 GeV. At higher centre-of-mass energies, other processes can play an important role such as the production in top quark decays via  $t \rightarrow b + H^\pm$  if  $m_{H^\pm}^2 < m_t - m_b$  or via the one-loop process  $e^+e^- \rightarrow W^\pm H^\mp$  [439, 440], which allows the production of a charged Higgs boson with  $m_{H^\pm}^2 > \sqrt{s}/2$ , even when  $H^+H^-$  production is kinematically forbidden. Other single charged Higgs production mechanisms include  $t\bar{b}H^-/\bar{t}bH^+$  production [107],  $\tau^+\nu H^-/\tau^-\bar{\nu}H^+$  production [441], and a variety of processes in which  $H^\pm$  is produced in association with a one or two other gauge and/or Higgs bosons [442].

At hadron colliders, Charged Higgs bosons can be produced in several different modes. If  $m_{H^\pm} < m_t - m_b$ , the charged Higgs boson can be produced in decays of the top quark via the decay  $t \rightarrow bH^\pm$ . Relevant QCD and SUSY-QCD corrections to  $\text{BR}(t \rightarrow H^\pm b)$  have been computed [299–302]. For values of  $m_{H^\pm}$  near  $m_t$ , width effects

are important. In addition, the full  $2 \rightarrow 3$  processes  $pp/p\bar{p} \rightarrow H^\pm \bar{t}b + X$  and  $pp/p\bar{p} \rightarrow H^\mp \bar{t}b + X$  must be considered. If  $m_{H^\pm} > m_t - m_b$ , then charged Higgs boson production occurs mainly through radiation from a third generation quark. Charged Higgs bosons may also be produced singly in association with a top quark via the  $2 \rightarrow 3$  partonic processes  $gq, q\bar{q} \rightarrow t\bar{b}H^-$ . For charged Higgs boson production cross section predictions for the Tevatron and the LHC, see Refs. [11, 44, 43]. Charged Higgs bosons can also be produced via associated production with  $W^\pm$  bosons through  $b\bar{b}$  annihilation and  $gg$ -fusion [303] and in pairs via  $q\bar{q}$  annihilation [304].

#### (a) Exclusion limits from LEP

Charged Higgs bosons have been searched for at LEP, where the combined data of the four experiments, ALEPH, DELPHI, L3, and OPAL, were sensitive to masses of up to about 90 GeV [421] in two decay channels, the  $\tau\nu$  and  $c\bar{s}$ . The exclusion limit independent of the admixture of the two above mentioned branching fractions was 78.6 GeV.

#### (b) Exclusion limits from Tevatron

Compared to the mass domain covered by LEP searches, the Tevatron covered a complementary range of charged Higgs masses. The CDF and D0 collaborations have also searched for charged Higgs bosons in top quark decays with subsequent decays to  $\tau\nu$  or to  $c\bar{s}$  [443–445]. For the  $H^+ \rightarrow c\bar{s}$  channel, the limits on  $\text{BR}(t \rightarrow H^+ b)$  from CDF and D0 are  $\approx 20\%$  in the mass range  $90 \text{ GeV} < m_{H^+} < 160 \text{ GeV}$  and assuming a branching fraction of 100% in this specific final state.  $H^+ \rightarrow \tau^+\nu_\tau$  channel, D0's limits on  $\text{BR}(t \rightarrow H^+ b)$  are also  $\approx 20\%$  in the same mass range and assuming a branching fraction of 100% in this final state. These limits are valid in general 2HDMs, and they have also been interpreted in terms of the MSSM [443–445].

#### (c) Exclusion limits from LHC

Similarly to the Tevatron, at the LHC light charged Higgs bosons can be searched for in the decays of top quarks. The main initial production mode for light charged Higgs bosons ( $m_{H^\pm} < m_t - m_b$ ) is top pair production. The subsequent decay modes of the charged Higgs boson for these searches are  $\tau\nu$  and  $c\bar{s}$ . More recently ATLAS and CMS have also searched for higher mass charged Higgs bosons ( $m_{H^\pm} > m_t + m_b$ ) in  $H^+ \rightarrow t\bar{b}$ . The main production modes are the associated production of a charged Higgs boson in association with a top and a bottom quark or in association with a top quark only.

ATLAS has searched for the decay  $H^+ \rightarrow \tau^+\nu_\tau$  in three final state topologies [446]: (i) lepton+jets: with  $t\bar{t} \rightarrow \bar{b}WH^+ \rightarrow b\bar{b}(q\bar{q}')(\tau_{\text{lep}}\nu)$ , *i.e.*, the  $W$  boson decays hadronically and the tau decays into an electron or a muon, with two neutrinos; (ii)  $\tau$  +lepton: with  $t\bar{t} \rightarrow \bar{b}WH^+ \rightarrow b\bar{b}(\ell\nu)(\tau_{\text{had}}\nu)$  *i.e.*, the  $W$  boson decays leptonically (with  $\ell = e, \mu$ ) and the tau decays hadronically; (iii)  $\tau$  +jets:  $t\bar{t} \rightarrow \bar{b}WH^+ \rightarrow b\bar{b}(q\bar{q}')(\tau_{\text{had}}\nu)$ , *i.e.*, both the  $W$  boson and the  $\tau$  decay hadronically [447]. Assuming  $\text{BR}(H^+ \rightarrow \tau^+\nu_\tau) = 100\%$ , ATLAS sets upper limits on  $\text{BR}(t \rightarrow H^+ b)$  between 0.24% and 2.1% for charged Higgs boson masses between 90 GeV to 160 GeV. When interpreted in the context of the  $m_h^{\text{max}}$  scenario of the MSSM, these bounds exclude a large fraction of the  $(m_{H^\pm}, \tan \beta)$  plane.

The CMS collaboration has searched for the charged Higgs boson in the decay products of top quark pairs:  $t\bar{t} \rightarrow H^\pm W^\mp b\bar{b}$  and  $t\bar{t} \rightarrow H^+H^-b\bar{b}$  [448, 449] as well. Three types of final states with large missing transverse energy and jets originating from  $b$ -quark hadronization have been analyzed: the fully-hadronic channel with a hadronically decaying tau in association with jets, the dilepton channel with a hadronically decaying tau in association with an electron or muon and the dilepton channel with an electron-muon pair. Combining the results of these three analyses and assuming  $\text{BR}(H^\pm \rightarrow \tau\nu) = 1$ , the upper limits on  $\text{BR}(t \rightarrow H^\pm b)$  are less than 2% to 3% depending on the charged Higgs boson mass in the interval  $80 \text{ GeV} < m_{H^\pm} < 160 \text{ GeV}$ .

Both the ATLAS [450] and CMS [449] experiments have also searched for high mass charged Higgs bosons decaying to a top and bottom quarks. The main production mode for this search is the

associated production with one top quark (5-flavour scheme) or a top quark and a bottom quark (4-flavour scheme) in the final state. The  $s$ -channel production mode where the charged Higgs boson is produced alone in the final state at tree level is also considered. This search is particularly intricate and it is sensitive to the modeling of the top pair production background produced in association with additional partons and in particular  $b$ -quarks. No excess was found and the results are expressed in terms of exclusion limits of cross section times branching fractions. The CMS collaboration has combined the results of this search with the  $H^+ \rightarrow \tau^+ \nu_\tau$  search in the framework of an updated mHmax MSSM scenario [449].

ATLAS and CMS have also searched for charged Higgs bosons in top quark decays assuming  $\text{BR}(H^+ \rightarrow c\bar{s}) = 100\%$  [451, 452], and sets limits of  $\approx 20\%$  on  $\text{BR}(t \rightarrow H^+ b)$  in the  $90 \text{ GeV} < m_{H^+} < 160 \text{ GeV}$  mass range.

In two Higgs doublet models the decay of the Charged Higgs boson to a  $W$ - and a  $Z$ -boson is allowed only at loop level and is therefore suppressed. However the  $H^\pm \rightarrow W^\pm Z$  decay channel is allowed in Higgs triplet models. The ATLAS experiment [453] has searched for such decays, requiring that the Charged Higgs boson is produced through the fusion of vector bosons. No excess with respect to the SM backgrounds has been observed in this channel, and the results are interpreted in the Georgi–Machacek model [360–362] discussed in Section VII.5.2.

At the LHC various other channels still remain to be explored, in particular searches involving additional neutral scalars in particular in  $WH$ ,  $WA$  where  $A$  is the pseudo scalar MSSM Higgs boson, and  $Wa$  where  $a$  is the light CP-odd scalars of the NMSSM.

#### (v) Searches for a light CP-odd Higgs boson $a$

A light pseudoscalar boson  $a$  is present in any two Higgs doublet mode enhanced with an additional singlet field. A prominent example is the NMSSM. The theoretical motivations for singlet extensions of the MSSM are discussed in Section VII.2. In the NMSSM, the searches now focus on the low mass pseudo-scalar boson  $a$  region for several reasons: (i) in the NMSSM, the light pseudo-scalar  $a$  boson can, as a pseudo-Goldstone boson, be a natural candidate for an axion; (ii) scenarios where  $m_a > 2m_b$  and a CP-even state  $h$  decaying to a pair of  $a$  ( $m_h > 2m_a$ ) are excluded by direct searches at LEP in the four  $b$  channel [421, 455, 456]; (iii) in the pre-discovery era, LEP limits on a CP-even Higgs boson resulted in fine tuning MSSM constraints [457] which could be evaded through non standard decays of the Higgs to  $aa$ ; (iv) an NMSSM CP-odd  $a$  boson with a mass in the range 9.2–12 GeV can also account for the difference observed between the measured anomalous muon magnetic moment and its prediction [458]. A scenario that has drawn particular attention was motivated by a small excess of events  $2.3\sigma$  in the SM Higgs search at LEP at Higgs boson mass of around 98 GeV. Speculative interpretations of this excess as a signal of a Higgs boson with reduced couplings to  $b$ -quarks were given [457]. Complete reviews of the NMSSM phenomenology can be found in Refs. [459, 456].

The potential benchmark scenarios have changed in the light of the  $H$  discovery. The discovered state could be the lightest or the next-to-lightest of the three CP-even states of the NMSSM. Light pseudoscalar scenarios are still very interesting in particular for the potential axion candidate. There are three main types of direct searches for the light  $a$  boson: (i) for masses below the  $\Upsilon$  resonance, the search is for radiative decays  $\Upsilon \rightarrow a\gamma$  at B-factories; (ii) the inclusive search for in high energy  $pp$  collisions at the LHC; (iii) the search for decays of a CP-even Higgs  $h$  boson to a pair of  $a$  bosons.

Radiative decays  $\Upsilon \rightarrow a\gamma$ , have been searched for in various colliders, the most recent results are searches for radiative decays of the  $\Upsilon(1s)$  to  $a\gamma$  with a subsequent decay of the  $a$  boson to a pair of taus at CLEO [460] and the radiative decays of the  $\Upsilon(1s, 2s, 3s)$  to  $a\gamma$  with subsequent decays to a pair of muons or taus by the BaBar collaboration [461, 462].

Direct inclusive searches for the light pseudo scalar  $a$  boson were performed in the  $a \rightarrow \mu\mu$  channel at the Tevatron by the D0 experiment [463] and by the ATLAS [464] and CMS [465] collaborations at the LHC.

Finally searches for the decays of the Higgs boson to a pair of  $a$  bosons were performed with subsequent decays to four photons by the ATLAS experiment [466], in the four muons final state by the CMS and D0 experiments [456, 467], in the two muons and two taus final state by the ATLAS [466] and D0 [456] collaborations, and in the four taus final state by the CMS [468] ALEPH collaboration at LEP [469].

No significant excess in the searches for a light CP-odd  $a$  boson were found and limits on the production times branching fractions of the  $a$  boson have been set.

#### (vi) Searches for doubly charged Higgs bosons $H^{\pm\pm}$

As discussed in Section VII.5, the generation of small neutrino masses via the standard EWSB mechanism described in Section II requires unnaturally small Yukawa couplings, provided that neutrinos are Dirac-type fermions. A Majorana mass term with a see-saw mechanism for neutrinos, would allow for naturally small masses and yield a framework for the appealing scenario of leptogenesis. However within the SM Majorana mass terms correspond to (non-renormalizable) dimension-5 operators. Such effective interactions can be generated via renormalizable interactions with an electroweak triplet of complex scalar fields (corresponding to a type-II see-saw mechanism). Other models such as the Zee–Babu model, with the introduction of two  $\text{SU}(2)_L$  singlets, also generate Majorana mass terms. The signature of such models would be the presence of doubly charged Higgs bosons  $H^{\pm\pm}$ .

The main production mechanisms of  $H^{\pm\pm}$  bosons at hadron colliders are the pair production in the  $s$ -channel through the exchange of a  $Z$  boson or a photon and the associated production with a Charged Higgs boson through the exchange of a  $W$  boson.

#### VII.8.1. Searches for non-standard production processes of the Higgs boson

The discovery of the Higgs boson has also allowed for searches of BSM (beyond the SM) processes involving standard decays of the Higgs boson. One example directly pertaining to the search for additional states of the EWSB sector is the search for Higgs bosons in the cascade decay of a heavy CP-even Higgs boson decaying to charged Higgs boson and a  $W$  boson, and the charged Higgs boson subsequently decaying to  $H$  and another  $W$  boson. This search has been performed by the ATLAS collaboration in  $b\bar{b}$  decays of the  $H$  particle [485].

Another example of searches for non standard processes through the presence of the  $H$  particle is the search for large flavor changing neutral current decays of the top quark to  $H$  and a charm quark. This search has been performed with the ATLAS experiment in the  $H \rightarrow \gamma\gamma$  channel [486].

#### VII.8.2. Outlook of searches for additional states

The LHC program of searches for additional states covers a large variety of decay and production channels. Since the last review on the *Status of Higgs boson physics* [243] many new channels have been explored at the LHC, *e.g.* the searches for additional states decaying into  $hh$  or  $Zh$ . The search for charged Higgs bosons has been extended to include the  $WZ$  and the very difficult  $t\bar{b}$  decay channel. There are however more channels to cover, *e.g.* the search for charged Higgs bosons in the  $HW$  and  $AW$  channels.

### VIII. Summary and outlook

*Summary*– The discovery of the Higgs boson is an important milestone in the history of particle physics as well as an extraordinary success of the LHC machine and the ATLAS and CMS experiments. Since its discovery, substantial progress in the field of Higgs boson physics has been accomplished and a significant number of measurements probing its nature have been made. They are revealing an increasingly precise profile of the Higgs boson. All experimental measurements are consistent with the EWSB mechanism of the Standard Model (SM).

Since the last review [487], the ATLAS and CMS experiments have made a combined measurement of the mass of the Higgs boson

in the diphoton and the four-lepton channels at per mille precision,  $m_H = 125.09 \pm 0.24$  GeV. The quantum numbers of the Higgs boson have been probed in greater detail and show an excellent consistency with the  $J^{PC} = 0^{++}$  hypothesis. Anomalous CP-even and CP-odd couplings have also been probed, mostly using angular distributions in diboson events. Higgs boson production and decay mechanisms have been further characterized through the measurement of various differential and fiducial cross sections.

The ATLAS and CMS collaborations have produced a detailed combined measurement of the properties of the Higgs boson. This combination establishes the direct observation of the VBF production process with a significance of  $5.4\sigma$  and the observation of the  $H \rightarrow \tau^+\tau^-$  with a significance of  $5.5\sigma$ . This combination also provides the most precise probes of the coupling structure of the Higgs boson, with a 10%-20% accuracy. This precision could not have been reached without the rapid and profound theoretical developments on many fronts: higher order calculations, Monte Carlo simulations, and new ideas on how to extract further informations on the nature of the Higgs boson. A particularly important breakthrough has been the recent calculation at the next-to-next-to-next-to-leading order of the Higgs production by gluon fusion, the dominant production channel at the LHC. All measurements are consistent with the SM predictions and provide stringent constraints on a large number of scenarios of new physics predicting sizeable deviations in the couplings of the Higgs boson.

Without assumptions on or a measurement of the Higgs boson width, the measurements at the LHC do not provide constraints on the absolute couplings of the Higgs boson. In the SM, the total Higgs width is approximately 4.2 MeV. The direct experimental measurements using the Higgs boson mass lineshape yield an upper bound still three orders of magnitude above its SM value. However, new ideas have emerged through the study of the Higgs couplings away from its mass shell. Under the specific assumption that the running of the Higgs couplings to vector bosons and gluons is small, the total width of the Higgs boson is constrained to be smaller than 25 MeV at 95% CL. Another interesting new idea, in the diphoton channel, is the observation that the interference between the signal and the continuum background induces shifts in the mass. Constraints on the width of the Higgs boson can then be inferred from precise measurements of its mass.

Further useful information on the components of the width of the Higgs boson can also be obtained from searches for rare and exotic decay modes, including invisible decays. Insights on the couplings of the Higgs boson are also obtained from the searches for rare production modes. No significant deviation from the SM Higgs boson expectations has been found in the channels analyzed so far.

Finally, all extensions of the SM at higher energies call for an enlargement of the EWSB sector. Therefore invaluable insights can also be acquired from searches for new additional scalar states. Since the last review [487], an ample number of new searches for CP-even and CP-Odd neutral Higgs bosons, and charged Higgs bosons have been carried out. No significant deviations from the minimal SM Higgs sector has been found in the ranges of mass and couplings of the additional states that have been explored so far.

**Outlook**– The unitarization of the vector boson scattering (VBS) amplitudes, dominated at high energies by their longitudinal polarizations, has been the basis of the *no lose* theorem at the LHC and was one of the main motivations to build the accelerator and the detectors. It motivated the existence of a Higgs boson or the observability of manifestations of strong dynamics at TeV scale. Now that a Higgs boson has been found and that its couplings to gauge bosons comply with the SM predictions, perturbative unitarity is preserved to a large amount with the sole exchange of the Higgs boson and without the need for any additional states. It is, however, still an important channel to investigate further in order to better understand the nature of the Higgs sector and the possible completion of the SM at the TeV scale. In association with the double Higgs boson production channel by vector boson fusion, VBS could, for instance, confirm that the Higgs boson is part of a weak doublet and also establish whether it is an elementary object or a composite state that could emerge as a pseudo-Nambu–Goldstone boson from a new

underlying broken symmetry.

The Higgs boson couplings are not dictated by any local gauge symmetry. Thus, in addition to a new particle, the LHC has also discovered a new force, different in nature from the other fundamental interactions since it is non-universal and distinguishes between the three families of quarks and leptons. The existence of the Higgs boson embodies the problem of an unnatural cancellation among the quantum corrections to its mass, if new physics is present at scale significantly higher than the EW scale. The non-observation of additional states which could stabilize the Higgs mass is a challenge for natural scenarios like supersymmetry or models with a new strong interaction in which the Higgs boson is not a fundamental particle. This increasingly pressing paradox starts questioning the principle of naturalness which underlies the hypothesis that phenomena at different scales do not influence each other.

The search for the Higgs boson has occupied the Particle physics community for the last 50 years. Its discovery has shaped and sharpened the physics programs of the LHC and of future accelerators. The experimental data together with the progress in theory mark the beginning of a new era of precision Higgs boson measurements.

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## References

1. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B716**, 1 (2012).
2. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B716**, 30 (2012).
3. S.L. Glashow, Nucl. Phys. **20**, 579 (1961);  
S. Weinberg, Phys. Rev. Lett. **19**, 1264 (1967);  
A. Salam, *Elementary Particle Theory*, eds.: Svartholm, Almqvist and Wiksells, Stockholm, 1968;  
S. Glashow, J. Iliopoulos, and L. Maiani, Phys. Rev. **D2**, 1285 (1970).
4. F. Englert and R. Brout, Phys. Rev. Lett. **13**, 321 (1964);  
P.W. Higgs, Phys. Rev. Lett. **13**, 508 (1964) and Phys. Rev. **145**, 1156 (1966);  
G.S. Guralnik, C.R. Hagen, and T.W. Kibble, Phys. Rev. Lett. **13**, 585 (1964).
5. J.M. Cornwall, D.N. Levin, and G. Tiktopoulos, Phys. Rev. Lett. **30**, 1286 (1973) and Phys. Rev. **D10**, 1145 (1974);  
C.H. Llewellyn Smith, Phys. Lett. **B46**, 233 (1973).
6. B.W. Lee, C. Quigg, and H.B. Thacker, Phys. Rev. **D16**, 1519 (1977).
7. K. Wilson, Phys. Rev. **D 3**, 1818 (1971);  
G. 't Hooft, in *Proc. of 1979 Cargèse Institute on Recent Developments in Gauge Theories*, p. 135 Press, New York 1980;  
For a recent review, see G.F. Giudice, PoS EPS **-HEP2013**, 163 (2013).



8. J. Parsons, and A. Pomarol, *Extra Dimensions*, in this volume.
9. J. Wess and B. Zumino, Nucl. Phys. **B70**, 39 (1974) and Phys. Lett. **49B**, 52 (1974);  
H.P. Nilles, Phys. Rev. **C110**, 1 (1984);  
S.P. Martin, [arXiv:hep-ph/9709356](#) (1997);  
P. Fayet, Phys. Lett. **B69**, 489 (1977), Phys. Lett. **B84**, 421 (1979), Phys. Lett. **B86**, 272 (1979) and Nucl. Phys. **B101**, 81 (2001).
10. E. Witten, Nucl. Phys. **B188**, 513 (1981);  
R.K. Kaul, Phys. Lett. **B109**, 19 (1982) and Pramana **19**, 183 (1982);  
L. Susskind, Phys. Rev. **104**, 181 (1984).
11. H.E. Haber and G.L. Kane, Phys. Rev. **C117**, 75 (1985).
12. J. F. Gunion *et al.*, *The Higgs Hunter's Guide*, Addison-Wesley (1990).
13. S. Weinberg, Phys. Rev. **D13**, 974 (1979) and Phys. Rev. **D19**, 1277 (1979);  
L. Susskind, Phys. Rev. **D20**, 2619 (1979);  
for a recent review, see C.T. Hill and E.H. Simmons, Phys. Reports **381**, 235 (2003) [E: **390**, 553 (2004)].
14. D. B. Kaplan and H. Georgi, Phys. Lett. **B136**, 183 (1984).
15. G. Panico and A. Wulzer, Lect. Notes Phys. **913**, 1 (2016).
16. C. Csaki, C. Grojean and J. Terning, [arXiv:1512.00468 \[hep-ph\]](#) (2015).
17. C. Csaki *et al.*, Phys. Rev. **D69**, 055006 (2004);  
C. Csaki *et al.*, Phys. Rev. Lett. **92**, 101802 (2004).
18. C. Csaki *et al.*, Phys. Rev. **D62**, 045015 (2000);  
for other references, see the 2014 edition of this review.
19. W. D. Goldberger, B. Grinstein, and W. Skiba, Phys. Rev. Lett. **100**, 111802 (2008);  
J. Fan *et al.*, Phys. Rev. **D79**, 035017 (2009);  
B. Bellazzini *et al.*, Eur. Phys. J. **C73**, 2333 (2013);  
Z. Chacko and R.K. Mishra, Phys. Rev. **D87**, 115006 (2013);  
Z. Chacko, R. Franceschini, and R.K. Mishra, JHEP **1304**, 015 (2013).
20. Z. Chacko, H.S. Goh and R. Harnik, Phys. Rev. Lett. **96**, 231802 (2006) and JHEP **0601**, 108 (2006).
21. N. Craig, A. Katz, M. Strassler and R. Sundrum, JHEP **1507**, 105 (2015);  
N. Craig, S. Knapen and P. Longhi, Phys. Rev. Lett. **114**, 061803 (2015).
22. P.W. Graham, D. E. Kaplan and S. Rajendran, Phys. Rev. Lett. **115**, 221801 (2015);  
J.R. Espinosa *et al.*, Phys. Rev. Lett. **115**, 251803 (2015).
23. T. van Ritbergen and R.G. Stuart, Phys. Rev. Lett. **82**, 488 (1999) and Nucl. Phys. **B564**, 343 (2000);  
M. Steinhauser and T. Seidensticker, Phys. Lett. **B467**, 271 (1999);  
D.M. Webber *et al.*, [MuLan Collab.], Phys. Rev. Lett. **106**, 041803 (2011).
24. J. Ellis, M.K. Gaillard, and D.V. Nanopoulos, Nucl. Phys. **B106**, 292 (1976).
25. M.A. Shifman *et al.*, Sov. J. Nucl. Phys. **30**, 711 (1979) [Yad. Fiz. **30**, 1368 (1979)].
26. P. Sikivie *et al.*, Nucl. Phys. **B173**, 189 (1980);  
H. Georgi, Ann. Rev. Nucl. and Part. Sci. **43**, 209 (1993).
27. M.J.G. Veltman, Nucl. Phys. **B123**, 89 (1977).
28. R.S. Chivukula *et al.*, Ann. Rev. Nucl. and Part. Sci. **45**, 255 (1995).
29. S. Willenbrock, [arXiv:0410370 \[hep-ph\]](#) (2004).
30. I.V. Krive and A.D. Linde, Nucl. Phys. **B117**, 265 (1976);  
for other references, see the 2014 edition of this review.
31. J. Elias-Miro *et al.*, Phys. Lett. **B709**, 222 (2012);  
G. Degrandi *et al.*, JHEP **1208**, 098 (2012);  
D. Buttazzo *et al.*, JHEP **1312**, 089 (2013);  
J.R. Espinosa *et al.*, JHEP **1509**, 174 (2015).
32. J.R. Espinosa and M. Quiros, Phys. Lett. **B353**, 257 (1995);  
G. Isidori, G. Ridolfi, and A. Strumia, Nucl. Phys. **B609**, 387 (2001).
33. V. Branchina, E. Messina and M. Sher, Phys. Rev. **D91**, 013003 (2015).
34. A. Hook *et al.*, JHEP **1501**, 061 (2015);  
J. Kearney, H. Yoo and K.M. Zurek, Phys. Rev. **D91**, 123537 (2015).
35. C.D. Froggatt and H.B. Nielsen, Phys. Lett. **B368**, 96 (1996);  
M. Shaposhnikov and C. Wetterich, Phys. Lett. **B683**, 196 (2010);  
M. Holthausen, K.S. Lim, and M. Lindner, JHEP **1202**, 037 (2012).
36. F.L. Bezrukov and M. Shaposhnikov, Phys. Lett. **B659**, 703 (2008);  
F.L. Bezrukov, A. Magnin, and M. Shaposhnikov, Phys. Lett. **B675**, 88 (2009).
37. A. Salvio, Phys. Lett. **B727**, 234 (2013).
38. B.A. Kniehl, Phys. Reports **240**, 211 (1994).
39. M. Spira, Fortsch. Phys. **46**, 203 (1998).
40. M. Carena and H.E. Haber, Prog. in Part. Nucl. Phys. **50**, 152 (2003).
41. A. Djouadi, Phys. Reports **457**, 1 (2008).
42. S. Dittmaier *et al.*, [LHC Higgs Cross Section Working Group], [arXiv:1101.0593 \[hep-ph\]](#) (2011).
43. S. Dittmaier *et al.*, [LHC Higgs Cross Section Working Group], [arXiv:1201.3084 \[hep-ph\]](#) (2012).
44. S. Heinemeyer *et al.*, [LHC Higgs Cross Section Working Group], [arXiv:1307.1347 \[hep-ph\]](#) (2013).
45. C. Anastasiou *et al.*, [LHC Higgs Cross Section Working Group], CERN Report, to appear (2016).
46. LHC Higgs Cross Section Working Group,  
<https://twiki.cern.ch/twiki/bin/view/LHCPhysics/LHCHSWG>.
47. T. Aaltonen *et al.*, [CDF and D0 Collaborations], Phys. Rev. **D88**, 052014 (2013).
48. H.M. Georgi *et al.*, Phys. Rev. Lett. **40**, 692 (1978).
49. D. Graudenz, M. Spira, and P.M. Zerwas, Phys. Rev. Lett. **70**, 1372 (1993).
50. M. Spira *et al.*, Nucl. Phys. **B453**, 17 (1995).
51. S. Dawson, Nucl. Phys. **B359**, 283 (1991);  
A. Djouadi, M. Spira, and P.M. Zerwas, Phys. Lett. **B264**, 440 (1991).
52. R.V. Harlander and W.B. Kilgore, Phys. Rev. Lett. **88**, 201801 (2002);  
C. Anastasiou and K. Melnikov, Nucl. Phys. **B646**, 220 (2002);  
V. Ravindran, J. Smith, and W.L. van Neerven, Nucl. Phys. **B665**, 325 (2003).
53. C. Anastasiou *et al.*, Phys. Rev. Lett. **114**, 212001 (2015);  
C. Anastasiou *et al.*, JHEP **1605**, 058 (2016).
54. R.V. Harlander and K.J. Ozeren, JHEP **0911**, 088 (2009);  
A. Pak, M. Rogal, and M. Steinhauser, JHEP **1002**, 025 (2010).
55. A. Djouadi and P. Gambino, Phys. Rev. Lett. **73**, 2528 (1994);  
S. Actis *et al.*, Phys. Lett. **B670**, 12 (2008);  
U. Aglietti *et al.*, Phys. Lett. **B595**, 432 (2004);  
G. Degrandi and F. Maltoni, Phys. Lett. **B600**, 255 (2004).
56. C. Anastasiou, R. Boughezal, and F. Petriello, JHEP **0904**, 003 (2009).
57. S. Catani *et al.*, JHEP **0307**, 028 (2003);  
S. Moch and A. Vogt, Phys. Lett. **B631**, 48 (2005);  
E. Laenen and L. Magnea, Phys. Lett. **B632**, 270 (2006);  
A. Idilbi *et al.*, Phys. Rev. **D73**, 077501 (2006);  
V. Ravindran, Nucl. Phys. **B752**, 173 (2006);  
V. Ahrens *et al.*, Eur. Phys. J. **C62**, 333 (2009).
58. V. Ahrens *et al.*, Phys. Lett. **B698**, 271 (2011);  
D. de Florian and M. Grazzini, Phys. Lett. **B718**, 117 (2012);  
C. Anastasiou *et al.*, JHEP **1204**, 004 (2012).
59. M. Bonvini *et al.*, [arXiv:1603.08000 \[hep-ph\]](#) (2016).
60. J.C. Collins, D.E. Soper and G.F. Sterman, Nucl. Phys. **B250**, 199 (1985).
61. D. de Florian *et al.*, JHEP **1111**, 064 (2011);  
T. Becher and M. Neubert, Eur. Phys. J. **C71**, 1665 (2011);  
J.Y. Chiu *et al.*, JHEP **1205**, 084 (2012);  
J. Wang *et al.*, Phys. Rev. **D86**, 094026 (2012);  
T. Becher, M. Neubert and D. Wilhelm, JHEP **1305**, 110 (2013).



62. S. Catani and M. Grazzini, *Eur. Phys. J.* **C72**, 2132 (2012) [*E: C72*, 2132 (2012)].
63. W.Y. Keung and F.J. Petriello, *Phys. Rev.* **D80**, 01007 (2009); S. Buhler *et al.*, *JHEP* **1207**, 115 (2012).
64. H. Mantler and M. Wiesemann, *Eur. Phys. J.* **C73**, 2467 (2013).
65. M. Grazzini and H. Sargsyan, *JHEP* **1309**, 129 (2013).
66. R.V. Harlander, T. Neumann, K.J. Ozeren and M. Wiesemann, *JHEP* **1208**, 139 (2012); T. Neumann and M. Wiesemann, *JHEP* **1411**, 150 (2014).
67. R.V. Harlander, S. Liebler, and H. Mantler, *Comp. Phys. Comm.* **184**, 1605 (2013).
68. D. de Florian, M. Grazzini, and Z. Kunszt, *Phys. Rev. Lett.* **82**, 5209 (1999); C.J. Glosser and C. R. Schmidt, *JHEP* **0212**, 016 (2002); V. Ravindran, J. Smith and W. L. Van Neerven, *Nucl. Phys.* **B634**, 247 (2002); X. Liu and F. Petriello, *Phys. Rev.* **D87**, 014018 (2013).
69. J.M. Campbell, R.K. Ellis, and G. Zanderighi, *JHEP* **0610**, 028 (2006); J.M. Campbell, R.K. Ellis, and C. Williams, *Phys. Rev.* **D81**, 074023 (2010).
70. R. Boughezal *et al.*, *JHEP* **1306**, 072 (2013); R. Boughezal *et al.*, *Phys. Rev. Lett.* **115**, 082003 (2015); X. Chen *et al.*, *Phys. Lett.* **B740**, 147 (2015) and *arXiv:1604.04085 [hep-ph]* (2016).
71. C.F. Berger *et al.*, *JHEP* **1104**, 092 (2011); A. Banfi, G.P. Salam, and G. Zanderighi, *JHEP* **1206**, 159 (2012); T. Becher and M. Neubert, *JHEP* **1207**, 108 (2012); A. Banfi *et al.*, *Phys. Rev. Lett.* **109**, 202001 (2012); F.J. Tackmann, J.R. Walsh, and S. Zuberi, *Phys. Rev.* **D86**, 053011 (2012); T. Becher, M. Neubert, and L. Rothen, *JHEP* **1310**, 125 (2013).
72. A. Banfi *et al.*, *JHEP* **1604**, 049 (2016).
73. I. Moullet and I. W. Stewart, *JHEP* **1409**, 129 (2014).
74. M. Dührssen *et al.*, *Phys. Rev.* **D70**, 113009 (2004).
75. T. Han, G. Valencia, and S. Willenbrock, *Phys. Rev. Lett.* **69**, 3274 (1992); T. Figy, C. Oleari, and D. Zeppenfeld, *Phys. Rev.* **D68**, 073005 (2003); T. Figy and D. Zeppenfeld, *Phys. Lett.* **B591**, 297 (2004); E.L. Berger and J. Campbell, *Phys. Rev.* **D70**, 073011 (2004); M. Ciccolini, A. Denner, and S. Dittmaier, *Phys. Rev. Lett.* **99**, 161803 (2007); M. Ciccolini, A. Denner, and S. Dittmaier, *Phys. Rev.* **D77**, 103002 (2008); A. Denner, S. Dittmaier, and A. Muck, *HAWK*, <http://omnibus.uni-freiburg.de/~sd565/programs/hawk/hawk.html>; K. Arnold *et al.*, *VBFNLO*, *Comp. Phys. Comm.* **180**, 1661 (2009); M. Spira, *VV2H*, <http://people.web.psi.ch/spira/vv2h>; N. Adam *et al.*, *arXiv:0803.1154 [hep-ph]* (2008); T. Figy, S. Palmer, and G. Weiglein, *JHEP* **1202**, 105 (2012).
76. P. Nason and C. Oleari, *JHEP* **1002**, 037 (2010); S. Frixione, P. Torrielli, and M. Zaro, *Phys. Lett.* **B726**, 273 (2013); F. Maltoni, K. Mawatari, and M. Zaro, *Eur. Phys. J.* **C74**, 2710 (2014).
77. P. Bolzoni *et al.*, *Phys. Rev. Lett.* **105**, 011801 (2010); P. Bolzoni *et al.*, *Phys. Rev.* **D85**, 035002 (2012).
78. M. Cacciari *et al.*, *Phys. Rev. Lett.* **115**, 082002 (2015).
79. S.L. Glashow, D.V. Nanopoulos, and A. Yildiz, *Phys. Rev.* **D18**, 1724 (1978); T. Han and S. Willenbrock, *Phys. Lett.* **B273**, 167 (1991); T. Han, G. Valencia, and S. Willenbrock, *Phys. Rev. Lett.* **69**, 3274 (1992); H. Baer, B. Bailey, and J.F. Owens, *Phys. Rev.* **D47**, 2730 (1993); J. Ohnemus and W.J. Stirling, *Phys. Rev.* **D47**, 2722 (1993).
80. A. Stange, W. Marciano, and S. Willenbrock, *Phys. Rev.* **D49**, 1354 (1994).
81. A. Stange, W. Marciano, and S. Willenbrock, *Phys. Rev.* **D50**, 4491 (1994).
82. M.L. Ciccolini, S. Dittmaier, and M. Kramer, *Phys. Rev.* **D68**, 073003 (2003); A. Denner, S. Dittmaier, and S. Kalweit, *JHEP* **1203**, 075 (2012).
83. R. Hamberg, W.L. van Neerven, and T. Matsuura, *Nucl. Phys.* **B359**, 343 (1991).
84. O. Brein, A. Djouadi, and R. Harlander, *Phys. Lett.* **B579**, 149 (2004); L. Altenkamp *et al.*, *JHEP* **1302**, 078 (2013).
85. O. Brein *et al.*, *Eur. Phys. J.* **C72**, 1868 (2012).
86. O. Brein, R.V. Harlander, and T.J. Zirke, *Comp. Phys. Comm.* **184**, 998 (2013).
87. A. Denner *et al.*, *JHEP* **1203**, 075 (2012).
88. G. Ferrera, M. Grazzini, and F. Tramontano, *Phys. Rev. Lett.* **107**, 152003 (2011).
89. G. Ferrera, M. Grazzini and F. Tramontano, *Phys. Lett.* **B740**, 51 (2015).
90. J.M. Campbell, R.K. Ellis and C. Williams, *JHEP* **1606**, 179 (2016).
91. W. Astill *et al.*, *arXiv:1603.01620 [hep-ph]* (2016).
92. R. Raitio and W.W. Wada, *Phys. Rev.* **D19**, 941 (1979); J.N. Ng and P. Zakarauskas, *Nucl. Phys.* **B247**, 339 (1984); J.F. Gunion, *Phys. Lett.* **B261**, 510 (1991); W.J. Marciano and F.E. Paige, *Phys. Rev. Lett.* **66**, 2433 (1991).
93. W. Beenakker *et al.*, *Phys. Rev. Lett.* **87**, 201805 (2001); L. Reina and S. Dawson, *Phys. Rev. Lett.* **87**, 201804 (2001); S. Dawson *et al.*, *Phys. Rev.* **D67**, 071503 (2003); W. Beenakker *et al.*, *Nucl. Phys.* **B653**, 151 (2003).
94. R. Frederix *et al.*, *Phys. Lett.* **B701**, 427 (2011); M. Garzelli *et al.*, *Europhys. Lett.* **96**, 11001 (2011).
95. F. Demartin *et al.*, *Eur. Phys. J.* **C75**, 267 (2015).
96. K.A. Assamagan *et al.*, [Higgs Working Group, “Physics at TeV Colliders” workshop, Les Houches, 2003], *arXiv:hep-ph/0406152* (2004).
97. R.V. Harlander and W.B. Kilgore, *Phys. Rev.* **D68**, 013001 (2003); J. M. Campbell *et al.*, *Phys. Rev.* **D67**, 095002 (2003); S. Dawson *et al.*, *Phys. Rev. Lett.* **94**, 031802 (2005); S. Dittmaier, M. Kramer, and M. Spira, *Phys. Rev.* **D70**, 074010 (2004); S. Dawson *et al.*, *Phys. Rev.* **D69**, 074027 (2004).
98. W.J. Stirling and D.J. Summers, *Phys. Lett.* **B283**, 411 (1992); F. Maltoni *et al.*, *Phys. Rev.* **D64**, 094023 (2001).
99. S. Dawson, S. Dittmaier and M. Spira, *Phys. Rev.* **D58**, 115012 (1998).
100. D. de Florian and J. Mazzitelli, *Phys. Rev. Lett.* **111**, 201801 (2013).
101. S. Borowka *et al.*, *Phys. Rev. Lett.* **117**, 012001 (2016).
102. B.L. Ioffe and V.A. Khoze, *Sov. J. Nucl. Phys.* **9**, 50 (1978).
103. D.R.T. Jones and S. Petcov, *Phys. Lett.* **B84**, 440 (1979); R.N. Cahn and S. Dawson, *Phys. Lett.* **B136**, 196 (1984); G.L. Kane, W.W. Repko, and W.B. Rolnick, *Phys. Lett.* **B148**, 367 (1984); G. Altarelli, B. Mele, and F. Pitolli, *Nucl. Phys.* **B287**, 205 (1987); W. Kilian, M. Kramer, and P.M. Zerwas, *Phys. Lett.* **B373**, 135 (1996).
104. B.A. Kniehl, *Z. Phys.* **C55**, 605 (1992).
105. J. Fleischer and F. Jegerlehner, *Nucl. Phys.* **B216**, 469 (1983); A. Denner *et al.*, *Z. Phys.* **C56**, 261 (1992).
106. B.A. Kniehl, *Int. J. Mod. Phys.* **A17**, 1457 (2002).
107. K.J. Gaemers and G.J. Gounaris, *Phys. Lett.* **B77**, 379 (1978); A. Djouadi, J. Kalinowski, and P. M. Zerwas, *Z. Phys.* **C54**, 255 (1992); B.A. Kniehl, F. Madricardo, and M. Steinhauser, *Phys. Rev.* **D66**, 054016 (2002).

108. S. Dittmaier *et al.*, Phys. Lett. **B441**, 383 (1998);  
S. Dittmaier *et al.*, Phys. Lett. **B478**, 247 (2000);  
S. Dawson and L. Reina, Phys. Rev. **D59**, 054012 (1999).
109. S. Dawson *et al.*, [Higgs Working Group, “Snowmass on the Mississippi” workshop] arXiv:1310.8361 [hep-ex] (2013).
110. D.M. Asner *et al.*, [ILC Higgs white paper, “Snowmass on the Mississippi” workshop] arXiv:1310.0763 [hep-ph] (2013).
111. A. Denner *et al.*, Eur. Phys. J. **C71**, 1753 (2011).
112. A. Djouadi, J. Kalinowski, and M. Spira, Comp. Phys. Comm. **108**, 56 (1998);  
A. Djouadi *et al.*, arXiv:1003.1643 [hep-ph] (2010).
113. S. Gorishnii *et al.*, Mod. Phys. Lett. **A5**, 2703 (1990);  
S. Gorishnii *et al.*, Phys. Rev. **D43**, 1633 (1991);  
A.L. Kataev and V.T. Kim, Mod. Phys. Lett. **A9**, 1309 (1994);  
L.R. Surguladze, Phys. Lett. **B341**, 60 (1994);  
S. Larin, T. van Ritbergen, and J. Vermaseren, Phys. Lett. **B362**, 134 (1995);  
K. Chetyrkin and A. Kwiatkowski, Nucl. Phys. **B461**, 3 (1996);  
K. Chetyrkin, Phys. Lett. **B390**, 309 (1997);  
P.A. Baikov, K.G. Chetyrkin, and J.H. Kuhn, Phys. Rev. Lett. **96**, 012003 (2006).
114. J. Fleischer and F. Jegerlehner, Phys. Rev. **D23**, 2001 (1981);  
D. Bardin, B. Vilenky, and P. Khristova, Sov. J. Nucl. Phys. **53**, 152 (1991);  
A. Dabelstein and W. Hollik, Z. Phys. **C53**, 507 (1992);  
B.A. Kniehl, Nucl. Phys. **B376**, 3 (1992);  
A. Djouadi *et al.*, *Proceedings  $e^+e^-$  collisions at 500 GeV* (1991).
115. T. Inami, T. Kubota, and Y. Okada, Z. Phys. **C18**, 69 (1983);  
K.G. Chetyrkin, B.A. Kniehl, and M. Steinhauser, Phys. Rev. Lett. **79**, 353 (1997);  
P.A. Baikov and K.G. Chetyrkin, Phys. Rev. Lett. **97**, 061803 (2006).
116. H.Q. Zheng and D.D. Wu, Phys. Rev. **D42**, 3760 (1990);  
A. Djouadi *et al.*, Phys. Lett. **B257**, 187 (1991);  
S. Dawson and R. Kauffman, Phys. Rev. **D47**, 1264 (1993);  
A. Djouadi, M. Spira, and P. Zerwas, Phys. Lett. **B311**, 255 (1993);  
K. Melnikov and O.I. Yakovlev, Phys. Lett. **B312**, 179 (1993);  
M. Inoue *et al.*, Mod. Phys. Lett. **A9**, 1189 (1994).
117. P. Maierhofer and P. Marquard, Phys. Lett. **B721**, 131 (2013).
118. U. Aglietti *et al.*, Phys. Lett. **B595**, 432 (2004);  
G. Degrandi and F. Maltoni, Phys. Lett. **B600**, 255 (2004);  
S. Actis *et al.*, Phys. Lett. **B670**, 12 (2008);  
U. Aglietti *et al.*, Phys. Lett. **B600**, 57 (2004);  
G. Degrandi and F. Maltoni, Nucl. Phys. **B724**, 183 (2005);  
U. Aglietti *et al.*, [Tevatron for LHC report: Higgs] arXiv:hep-ph/0612172 (2006).
119. A. Abbasabadi *et al.*, Phys. Rev. **D55**, 5647 (1997);  
A. Abbasabadi and W.W. Repko, Phys. Rev. **D71**, 017304 (2005);  
A. Abbasabadi and W.W. Repko, JHEP **0608**, 048 (2006);  
D.A. Dicus and W.W. Repko, Phys. Rev. **D87**, 077301 (2013);  
L.B. Chen, C.F. Qiao, and R.L. Zhu, Phys. Lett. **B726**, 306 (2013);  
Y. Sun, H.R. Chang, and D.N. Gao, JHEP **1305**, 061 (2013);  
G. Passarino, Phys. Lett. **B727**, 424 (2013).
120. M. Spira, A. Djouadi, and P.M. Zerwas, Phys. Lett. **B276**, 350 (1992).
121. A. Bredenstein *et al.*, Phys. Rev. **D74**, 013004 (2006);  
A. Bredenstein *et al.*, JHEP **0702**, 080 (2007);  
A. Bredenstein *et al.*, Prophecy4f: A Monte Carlo generator for a proper description of the Higgs decay into 4 fermions, <http://omnibus.uni-freiburg.de/~sd565/programs/prophecy4f/prophecy4f.html>.
122. A. Ghinculov, Phys. Lett. **B337**, 137 (1994) [E: **B346**, 426 (1995)];  
L. Durand, B.A. Kniehl, and K. Riesselmann, Phys. Rev. **D51**, 5007 (1995);  
L. Durand, K. Riesselmann, and B.A. Kniehl, Phys. Rev. Lett. **72**, 2534 (1994) [E: **74**, 1699 (1995)].
123. E. Braaten and J.P. Leveille, Phys. Rev. **D22**, 715 (1980);  
L. Durand, K. Riesselmann, and B.A. Kniehl, Phys. Rev. Lett. **72**, 2534 (1994);  
E. Gross, G. Wolf, and B.A. Kniehl, Z. Phys. **C63**, 417 (1994) [E: *ibid.*, **C66**, 32 (1995)];  
A. Ghinculov, Phys. Lett. **B337**, 137 (1994) and Nucl. Phys. **B455**, 21 (1995);  
A. Djouadi, M. Spira, and P.M. Zerwas, Z. Phys. **C70**, 427 (1996);  
A. Frink *et al.*, Phys. Rev. **D54**, 4548 (1996);  
K.G. Chetyrkin and M. Steinhauser, Phys. Lett. **B408**, 320 (1997);  
R. Harlander and M. Steinhauser, Phys. Rev. **D56**, 3980 (1997);  
A.L. Kataev, Sov. Phys. JETP Lett. **66**, 327 (1997);  
S. Actis *et al.*, Nucl. Phys. **B811**, 182 (2009).
124. J. Erler and A. Freitas, *Electroweak Model and Constraints on New Physics*, in this volume.
125. R. Barate *et al.*, [LEP Working Group for Higgs boson searches and ALEPH, DELPHI, L3, and OPAL Collaborations], Phys. Lett. **B565**, 61 (2003).
126. CDF and D0 Collaborations, Phys. Rev. **D88**, 052014 (2013).
127. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D90**, 112015 (2014).
128. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Rev. **D89**, 092007 (2014).
129. S. Chatrchyan *et al.*, [CMS Collab.], Eur. Phys. J. **C74**, 3076 (2014).
130. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D91**, 012006 (2015).
131. S. Chatrchyan *et al.*, [CMS Collab.], JHEP **12**, 034 (2012).
132. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D90**, 052004 (2014).
133. S. Chatrchyan *et al.*, [CMS Collab.], Eur. Phys. J. **C75**, 212 (2015).
134. G. Aad *et al.*, [Atlas and CMS Collaborations], Phys. Rev. Lett. **114**, 191803 (2015).
135. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D92**, 012006 (2015).
136. S. Chatrchyan *et al.*, [CMS Collab.], JHEP **01**, 096 (2014).
137. G. Aad *et al.*, [ATLAS Collab.], JHEP **08**, 137 (2015).
138. S. Chatrchyan *et al.*, [CMS Collab.], JHEP **05**, 104 (2014).
139. G. Aad *et al.*, [ATLAS Collab.], JHEP **04**, 117 (2015).
140. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Rev. **D89**, 012003 (2014).
141. ATLAS and CMS Collaborations, ATLAS-CONF-2015-044 and CMS-PAS-HIG-15-002 (2015).
142. T. Aaltonen *et al.*, [CDF and D0 Collaborations], Phys. Rev. Lett. **109**, 071804 (2012).
143. J.M. Butterworth *et al.*, Phys. Rev. Lett. **100**, 242001 (2008).
144. G. Aad *et al.*, [ATLAS Collab.], JHEP **01**, 069 (2015).
145. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Rev. **D92**, 032008 (2015).
146. ATLAS Collab., ATLAS-CONF-2015-060 (2015).
147. CMS Collab., CMS-PAS-HIG-15-005 (2016).
148. ATLAS Collab., ATLAS-CONF-2015-059 (2015).
149. CMS Collab., CMS-PAS-HIG-15-004 (2016).
150. ATLAS Collab., ATLAS-CONF-2015-069 (2015).
151. CMS Collab., CMS-PAS-HIG-15-003 (2016).
152. CMS Collab., CMS-PAS-HIG-16-003 (2016).
153. S. Chatrchyan *et al.*, [CMS Collab.], JHEP **09**, 087 (2014).
154. S. Chatrchyan *et al.*, [CMS Collab.], Eur. Phys. J. **C75**, 215 (2015).
155. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B740**, 222 (2012).
156. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C75**, 349 (2015).
157. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B749**, 519 (2015).
158. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B740**, 222 (2015).
159. G. Aad *et al.*, [ATLAS Collab.], JHEP **12**, 061 (2015).
160. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Rev. **D90**, 112013 (2014).
161. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. Lett. **114**, 081802 (2015).
162. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. **D92**, 092004 (2015).
163. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C75**, 412 (2015).

164. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Rev. **D90**, 112013 (2014).
165. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B755**, 220 (2016).
166. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B749**, 560 (2015).
167. ATLAS Collab., ATL-PHYS-PUB-2015-046 (2015).
168. CMS Collab., CMS-PAS-FTR-15-002 (2015).
169. ATLAS Collab., ATL-PHYS-PUB-2014-019 (2014).
170. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B726**, 587 (2013).
171. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B753**, 341 (2016).
172. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B738**, 68 (2014).
173. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B744**, 184 (2015).
174. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B749**, 137 (2015).
175. G. Aad *et al.*, [ATLAS Collab.], JHEP **1511**, 211 (2015).
176. G. Aad *et al.*, [ATLAS Collab.], CERN-EP-2016-055(2016).
177. C. Delaunay *et al.*, Phys. Rev. **D89**, 033014 (2014).
178. G. T. Bodwin *et al.*, Phys. Rev. **D88**, 053003 (2013).
179. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. Lett. **114**, 121801 (2015).
180. ATLAS Collab., CERN-EP-2016-130 (2016).
181. A. Djouadi *et al.*, Eur. Phys. J. **C73**, 2455 (2013).
182. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C75**, 337 (2015).
183. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. Lett. **112**, 201802 (2014).
184. G. Aad *et al.*, [ATLAS Collab.], JHEP **01**, 172 (2016).
185. G. Aad *et al.*, [ATLAS Collab.], JHEP **11**, 206 (2015).
186. S. Chatrchyan *et al.*, [CMS Collab.], Eur. Phys. J. **C74**, 2980 (2014).
187. S. Chatrchyan *et al.*, [CMS Collab.], CMS-PAS-HIG-15-012 (2015).
188. S. Chatrchyan *et al.*, [CMS Collab.], CMS-PAS-HIG-16-009 (2016).
189. S. Chatrchyan *et al.*, [CMS Collab.], CMS-PAS-HIG-16-008 (2016).
190. M.J. Strassler and K.M. Zurek, Phys. Lett. **B651**, 374 (2007).
191. M.J. Strassler and K.M. Zurek, Phys. Lett. **B661**, 263 (2008).
192. T. Han *et al.*, JHEP **0807**, 008 (2008).
193. A. Falkowski *et al.*, JHEP **1005**, 077 (2010) and Phys. Rev. Lett. **105**, 241801 (2010).
194. G. Aad *et al.*, [ATLAS Collab.], New J. Phys. **15**, 043009 (2013).
195. G. Aad *et al.*, [ATLAS Collab.], Phys. Lett. **B721**, 32 (2013).
196. G. Aad *et al.*, [ATLAS Collab.], Phys. Rev. Lett. **108**, 251801 (2012).
197. D. Tucker-Smith and N. Weiner, Phys. Rev. **D64**, 043502 (2001).
198. S. Chatrchyan *et al.*, [CMS Collab.], Phys. Lett. **B726**, 564 (2013).
199. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C76**, 6 (2016).
200. L.D. Landau, Dokl. Akad. Nauk Ser. Fiz. **60**, 207 (1948); C.N. Yang, Phys. Rev. **D77**, 242 (1950).
201. G. Aad *et al.*, [ATLAS Collab.], Eur. Phys. J. **C75**, 476 (2015).
202. V. Khachatryan *et al.*, [CMS Collab.], Phys. Rev. **D92**, 012004 (2015).
203. ATLAS Collab., Phys. Lett. **B726**, 120 (2013).
204. P. Artoisenet *et al.*, JHEP **1311**, 043 (2013); A. Alloul, B. Fuks, and V. Sanz, JHEP **1404**, 110 (2014); A. Falkowski *et al.*, Eur. Phys. J. **C75**, 583 (2015).
205. J. Ellis *et al.*, JHEP **1211**, 134 (2012).
206. D0 Collab., Note 6387-CONF (2013).
207. D0 Collab., Note 6406-CONF (2013).
208. A. De Rujula *et al.*, Phys. Rev. **D82**, 013003 (2010).
209. ATLAS Collab., Eur. Phys. J. **C75**, 335 (2015).
210. V. Khachatryan *et al.*, [CMS Collab.], Phys. Rev. **D92**, 072010 (2015).
211. V. Khachatryan *et al.*, [CMS Collab.], CERN-EP-2016-054.
212. L. Dixon and S. Siu, Phys. Rev. Lett. **90**, 252001 (2003); S. P. Martin, Phys. Rev. **D86**, 073016 (2012); L. Dixon and Y. Li, Phys. Rev. Lett. **111**, 111802 (2013).
213. ATLAS Collab., ATL-PHYS-PUB-2016-009 (2016).
214. ATLAS Collab., ATL-PHYS-PUB-2013-014 (2013).
215. N. Kauer and G. Passarino, JHEP **08**, 116 (2012); F. Caola and K. Melnikov, Phys. Rev. **D88**, 054024 (2013); J.M. Campbell, R.K. Ellis, and C. Williams, JHEP **04**, 060 (2014); J.M. Campbell, R.K. Ellis, and C. Williams, Phys. Rev. **D89**, 053011 (2014); C. Englert and M. Spannowsky, Phys. Rev. **D90**, 053003 (2014).
216. ATLAS Collab., ATL-PHYS-PUB-2015-024 (2015).
217. M. Gonzalez-Alonso *et al.*, Eur. Phys. J. **C75**, 128 (2015); A. Greljo *et al.*, Eur. Phys. J. **C76**, 158 (2016).
218. G. Buchalla, O. Cata and C. Krause, Nucl. Phys. **B880**, 552 (2014); I. Brivio *et al.*, JHEP **1403**, 014 (2014).
219. W. Buchmuller and D. Wyler, Nucl. Phys. **B268**, 621 (1986).
220. C.J. C. Burges and H.J. Schnitzer, Nucl. Phys. **B288**, 464 (1983); C.N. Leung, S.T. Love and S. Rao, Z. Phys. **C31**, 433 (1986).
221. B. Grzadkowski *et al.*, JHEP **1010**, 085 (2010).
222. L. Lehman and A. Martin, JHEP **1602**, 081 (2016); B. Henning *et al.*, arXiv:1512.03433 [hep-ph] (2015).
223. R. Alonso *et al.*, JHEP **1404**, 159 (2014).
224. G.F. Giudice *et al.*, JHEP **0706**, 045 (2007).
225. R. Contino *et al.*, JHEP **1307**, 035 (2013).
226. J. Elias-Miro *et al.*, JHEP **1311**, 066 (2013); R.S. Gupta, A. Pomarol and F. Riva, Phys. Rev. **D91**, 035001 (2015); E. Masso, JHEP **1410**, 128 (2014).
227. A. Falkowski and F. Riva, JHEP **1502**, 039 (2015).
228. A. Falkowski *et al.*, Phys. Rev. Lett. **116**, 011801 (2016).
229. C. Degrande *et al.*, JHEP **1207**, 036 (2012); J.F. Kamenik, M. Papucci, and A. Weiler, Phys. Rev. **D85**, 071501 (2012).
230. A. David *et al.*, [LHC Higgs Cross Section Working Group], arXiv:1209.0040 [hep-ph] (2012).
231. B.A. Kniehl and M. Spira, Z. Phys. **C69**, 77 (1995).
232. G. Isidori, A.V. Manohar, and M. Trott, Phys. Lett. **B728**, 131 (2014); G. Isidori and M. Trott, JHEP **1402**, 082 (2014).
233. A. Pomarol and F. Riva, JHEP **1401**, 151 (2014).
234. R. Godbole *et al.*, Phys. Lett. **B730**, 275 (2014).
235. M. Reece, New J. Phys. **15**, 043003 (2013).
236. S. Biswas, E. Gabrielli, and B. Mele, JHEP **1301**, 088 (2013); S. Biswas *et al.*, JHEP **07**, 073 (2013).
237. M. Farina *et al.*, JHEP **1305**, 022 (2013).
238. ATLAS Collab., JHEP **11**, 206 (2015).
239. A. Djouadi *et al.*, Eur. Phys. J. **C73**, 2455 (2013).
240. ATLAS Collab., ATLAS-CONF-2013-072 (2013).
241. ATLAS Collab., Phys. Lett. **B753**, 69 (2016).
242. L.E. Ibanez and G.G. Ross, Phys. Lett. **B110**, 215 (1982); L.E. Ibanez, Phys. Lett. **B118**, 73 (1982); J. Ellis, D.V. Nanopoulos, and K. Tamvakis, Phys. Lett. **B121**, 123 (1983); L. Alvarez-Gaume, J. Polchinski, and M.B. Wise, Nucl. Phys. **B221**, 495 (1983).
243. See the list of references in the corresponding section of the 2014 edition of this review.
244. ATLAS Collab., <https://twiki.cern.ch/twiki/bin/view/AtlasPublic/Publications>; CMS Collab., <http://cms-results.web.cern.ch/cms-results/public-results/publications/SUS/STOP.html>.
245. J. Mrazek *et al.*, Nucl. Phys. **B853**, 1 (2011).
246. H. Georgi and D. B. Kaplan, Phys. Lett. **B145**, 216 (1984).
247. D.B. Kaplan, Nucl. Phys. **B365**, 259 (1991).

248. E. Savioni, PhD thesis, <http://paduaresearch.cab.unipd.it/6166/> (2013).
249. G. Panico *et al.*, JHEP **1303**, 051 (2013).
250. H.E. Haber, *Supersymmetry*, in this volume.
251. A. Djouadi, Phys. Reports **459**, 1 (2008).
252. H.E. Haber and Y. Nir, Nucl. Phys. **B335**, 363 (1990);  
A. Dabelstein, Nucl. Phys. **B456**, 25 (1995);  
S. Heinemeyer, W. Hollik, and G. Weiglein, Eur. Phys. J. **C16**, 139 (2000);  
A. Dobado, M. J. Herrero, and S. Penaranda, Eur. Phys. J. **C17**, 487 (2000).
253. J.F. Gunion and H.E. Haber, Phys. Rev. **D67**, 075019 (2003).
254. N. Craig, J. Galloway, and S. Thomas, [arXiv:1305.2424](https://arxiv.org/abs/1305.2424) [hep-ph] (2013).
255. M. Carena *et al.*, JHEP **1203**, 014 (2012);  
M. Carena *et al.*, JHEP **1207**, 175 (2012).
256. G. Degrandi *et al.*, Eur. Phys. J. **C28**, 133 (2003).
257. S. Heinemeyer *et al.*, JHEP **0808**, 087 (2008).
258. M. Carena *et al.*, Eur. Phys. J. **C26**, 601 (2003);  
M. Carena *et al.*, Eur. Phys. J. **C73**, 2552 (2013).
259. M. Carena *et al.*, Eur. Phys. J. **C45**, 797 (2006).
260. S.Y. Choi, M. Drees, and J.S. Lee, Phys. Lett. **B481**, 57 (2000);  
M. Carena *et al.*, Nucl. Phys. **B625**, 345 (2002).
261. A. Pilaftsis and C.E.M. Wagner, Nucl. Phys. **B553**, 3 (1999).
262. A. Arbey *et al.*, Phys. Lett. **B708**, 162 (2012);  
A. Arbey *et al.*, JHEP **1209**, 107 (2012).
263. L.J. Hall, D. Pinner, and J.T. Ruderman, JHEP **1204**, 131 (2012).
264. H. Baer, V. Barger, and A. Mustafayev, Phys. Rev. **D85**, 075010 (2012).
265. P. Draper *et al.*, Phys. Rev. **D85**, 095007 (2012).
266. S. Heinemeyer, O. Stal, and G. Weiglein, Phys. Lett. **B710**, 201 (2012).
267. M. Kadastik *et al.*, JHEP **1205**, 061 (2012).
268. P. Draper, G. Lee and C.E.M. Wagner, Phys. Rev. **D89**, 055023 (2014).
269. J.P. Vega and G. Villadoro, JHEP **1507**, 159 (2015).
270. S.P. Martin, Phys. Rev. **D75**, 055005 (2007);  
P. Kant *et al.*, JHEP **1008**, 104 (2010);  
J.L. Feng *et al.*, Phys. Rev. Lett. **111**, 131802 (2013).
271. M. Carena *et al.*, Nucl. Phys. **B580**, 29 (2000).
272. T. Hahn *et al.*, Phys. Rev. Lett. **112**, 141801 (2014).
273. T.D. Lee, Phys. Rev. **D8**, 1226 (1973);  
P. Fayet, Nucl. Phys. **B78**, 14 (1974);  
R.D. Peccei and H.R. Quinn, Phys. Rev. Lett. **38**, 1440 (1977);  
P. Fayet and S. Ferrara, Phys. Reports **32**, 249 (1977);  
L.J. Hall and M.B. Wise, Nucl. Phys. **B187**, 397 (1981);  
V.D. Barger, J.L. Hewett, and R.J.N. Phillips, Phys. Rev. **D41**, 3421 (1990).
274. M. Carena, S. Mrenna, and C.E.M. Wagner, Phys. Rev. **D60**, 075010 (1999) and Phys. Rev. **D62**, 055008 (2000).
275. D.M. Pierce *et al.*, Nucl. Phys. **B491**, 3 (1997).
276. A. Dabelstein, Nucl. Phys. **B456**, 25 (1995);  
F. Borzumati *et al.*, Nucl. Phys. **B555**, 53 (1999);  
H. Eberl *et al.*, Phys. Rev. **D62**, 055006 (2000).
277. J.A. Coarasa Perez, R.A. Jimenez, and J. Sola, Phys. Lett. **B389**, 312 (1996);  
R.A. Jimenez and J. Sola, Phys. Lett. **B389**, 53 (1996);  
A. Bartl *et al.*, Phys. Lett. **B378**, 167 (1996).
278. S. Heinemeyer, W. Hollik, and G. Weiglein, Eur. Phys. J. **C16**, 139 (2000).
279. H. E. Haber *et al.*, Phys. Rev. **D63**, 055004 (2001).
280. L. Hall, R. Rattazzi, and U. Sarid, Phys. Rev. **D50**, 7048 (1994);  
R. Hempfling, Phys. Rev. **D49**, 6168 (1994).
281. M.S. Carena *et al.*, Nucl. Phys. **B426**, 269 (1994).
282. J. Guasch, P. Haffiger, and M. Spira, Phys. Rev. **D68**, 115001 (2003);  
D. Noth and M. Spira, Phys. Rev. Lett. **101**, 181801 (2008) and JHEP **1106**, 084 (2011);  
L. Mihaila and C. Reisser, JHEP **1008**, 021 (2010).
283. M.S. Carena *et al.*, Phys. Lett. **B499**, 141 (2001).
284. A. Djouadi, J. Kalinowski, and P.M. Zerwas, Z. Phys. **C57**, 569 (1993);  
H. Baer *et al.*, Phys. Rev. **D47**, 1062 (1993);  
A. Djouadi *et al.*, Phys. Lett. **B376**, 220 (1996);  
A. Djouadi *et al.*, Z. Phys. **C74**, 93 (1997);  
S. Heinemeyer and W. Hollik, Nucl. Phys. **B474**, 32 (1996).
285. J.F. Gunion, Phys. Rev. Lett. **72**, 199 (1994);  
D. Choudhury and D.P. Roy, Phys. Lett. **B322**, 368 (1994);  
O.J. Eboli and D. Zeppenfeld, Phys. Lett. **B495**, 147 (2000);  
B.P. Kersevan, M. Malawski, and E. Richter-Was, Eur. Phys. J. **C29**, 541 (2003).
286. E.L. Berger *et al.*, Phys. Rev. **D66**, 095001 (2002).
287. A. Brignole *et al.*, Nucl. Phys. **B643**, 79 (2002);  
R. Dermisek and I. Low, Phys. Rev. **D77**, 035012 (2008).
288. A. Djouadi, Phys. Lett. **B435**, 101 (1998).
289. M.R. Buckley and D. Hooper, Phys. Rev. **D86**, 075008 (2012);  
for other references, see the 2014 edition of this review.
290. J.J. Cao *et al.*, JHEP **1203**, 086 (2012);  
R. Benbrik *et al.*, Eur. Phys. J. **C72**, 2171 (2012);  
Z. Kang, J. Li, and T. Li, JHEP **1211**, 024 (2012);  
J.F. Gunion, Y. Jiang, and S. Kraml, Phys. Lett. **B710**, 454 (2012);  
U. Ellwanger, JHEP **1203**, 044 (2012).
291. LHC Cross Section Working Group for Beyond the Standard Model Higgs: NMSSM  
<https://twiki.cern.ch/twiki/bin/view/LHCPhysics/LHCHXSWGNMSSM>.
292. M. Carena *et al.*, Phys. Rev. **D93**, 035013 (2016).
293. T. Kitahara, JHEP **1211**, 021 (2012);  
M. Carena *et al.*, JHEP **1302**, 114 (2013).
294. B. Batell, S. Jung, and C.E.M. Wagner, JHEP **1312**, 075 (2013).
295. D. Dicus *et al.*, Phys. Rev. **D59**, 094016 (1999).
296. C. Balazs, H.-J. He, and C.P. Yuan, Phys. Rev. **D60**, 114001 (1999).
297. E. Boos *et al.*, Phys. Rev. **D66**, 055004 (2002);  
E. Boos, A. Djouadi, and A. Nikitenko, Phys. Lett. **B578**, 384 (2004);  
E. Boos *et al.*, Phys. Lett. **B622**, 311 (2005);  
M. Carena *et al.*, JHEP **1207**, 091 (2012).
298. A.A. Barrientos Bendezu and B.A. Kniehl, Phys. Rev. **D64**, 035006 (2001).
299. J.A. Coarasa Perez *et al.*, Eur. Phys. J. **C2**, 373 (1998);  
J.A. Coarasa Perez *et al.*, Phys. Lett. **B425**, 329 (1998).
300. C.S. Li and T.C. Yuan, Phys. Rev. **D42**, 3088 (1990) [E: **D47**, 2156 (1993)];  
A. Czarnecki and S. Davidson, Phys. Rev. **D47**, 3063 (1993);  
C.S. Li, Y.-S. Wei, and J.-M. Yang, Phys. Lett. **B285**, 137 (1992).
301. J. Guasch, R.A. Jimenez, and J. Sola, Phys. Lett. **B360**, 47 (1995).
302. M.S. Carena *et al.*, Nucl. Phys. **B577**, 88 (2000).
303. A.A. Barrientos Bendezu and B.A. Kniehl, Phys. Rev. **D59**, 015009 (1999), Phys. Rev. **D61**, 015009 (2000) and Phys. Rev. **D63**, 015009 (2001).
304. A.A. Barrientos Bendezu and B.A. Kniehl, Nucl. Phys. **B568**, 305 (2000).
305. A. Krause *et al.*, Nucl. Phys. **B519**, 85 (1998);  
O. Brein and W. Hollik, Eur. Phys. J. **C13**, 175 (2000).
306. R.M. Barnett, H.E. Haber, and D.E. Soper, Nucl. Phys. **B306**, 697 (1988);  
for other references, see the 2014 edition of this review.
307. L. Maiani, A. D. Polosa and V. Riquer, New J. Phys. **14**, 073029 (2012) and Phys. Lett. **B718**, 465 (2012);  
A. Djouadi and J. Quevillon, JHEP **1310**, 028 (2013);  
A. Djouadi *et al.*, Eur. Phys. J. **C73**, 2650 (2013).
308. M. Carena *et al.*, Phys. Rev. **D91**, 035003 (2015).
309. M. Carena *et al.*, JHEP **1404**, 015 (2014).
310. K. Blum, R.T. D'Agnolo, and J. Fan, JHEP **1301**, 057 (2013);  
A. Azatov *et al.*, Phys. Rev. **D86**, 075033 (2012);

- J.R. Espinosa *et al.*, JHEP **1212**, 077 (2012);  
 R.S. Gupta, M. Montull, and F. Riva, JHEP **1304**, 132 (2013);  
 R.T. D'Agnolo, PhD thesis, Scuola Normale Superiore, Pisa, 2013.
311. A. Djouadi *et al.*, JHEP **1506**, 168 (2015).  
 312. G. D'Ambrosio *et al.*, Nucl. Phys. **B645**, 155 (2002);  
 R.S. Chivukula and H. Georgi, Phys. Lett. **B188**, 99 (1987);  
 L.J. Hall and L. Randall, Phys. Rev. Lett. **65**, 2939 (1990);  
 A.J. Buras *et al.*, Phys. Lett. **B500**, 161 (2001).  
 313. L.J. Hall, J. Lykken, and S. Weinberg, Phys. Rev. **D27**, 2359 (1983);  
 J.E. Kim and H.P. Nilles, Phys. Lett. **B138**, 150 (1984);  
 G.F. Giudice and A. Masiero, Phys. Lett. **B206**, 480 (1988);  
 E.J. Chun, J.E. Kim, and H.P. Nilles, Nucl. Phys. **B370**, 105 (1992);  
 I. Antoniadis *et al.*, Nucl. Phys. **B432**, 187 (1994).  
 314. R.D. Peccei and H.R. Quinn, Phys. Rev. Lett. **38**, 1440 (1977).  
 315. P. Fayet, Phys. Lett. **B90**, 104 (1975).  
 316. Y.B. Zeldovich, I.Y. Kobzarev, and L.B. Okun, Zh. Eksp. Teor. Fiz. **67**, 3 (1974);  
 A. Vilenkin, Phys. Reports **121**, 263 (1985).  
 317. C. Panagiotakopoulos and K. Tamvakis, Phys. Lett. **B469**, 145 (1999);  
 A. Dedes *et al.*, Phys. Rev. **D63**, 055009 (2001);  
 A. Menon, D. Morrissey, and C.E.M. Wagner, Phys. Rev. **D70**, 035005 (2004).  
 318. M. Cvetič *et al.*, Phys. Rev. **D56**, 2861 (1997) [E: **D58**, 119905 (1998)];  
 P. Langacker and J. Wang, Phys. Rev. **D58**, 115010 (1998) and references therein.  
 319. J. Erler, P. Langacker, and T.J. Li, Phys. Rev. **D66**, 015002 (2002);  
 T. Han, P. Langacker and B. McElrath, Phys. Rev. **D70**, 115006 (2004);  
 V. Barger *et al.*, Phys. Rev. **D73**, 115010 (2006).  
 320. V. Barger *et al.*, Phys. Rev. **D73**, 115010 (2006);  
 V. Barger, P. Langacker, and G. Shaughnessy, Phys. Rev. **D75**, 055013 (2007).  
 321. E. Accomando, *et al.*, hep-ph/0608079 (2006).  
 322. B. A. Dobrescu, G. L. Landsberg, and K. T. Matchev, Phys. Rev. **D63**, 075003 (2001).  
 323. R. Dermisek and J. F. Gunion, Phys. Rev. Lett. **95**, 041801 (2005).  
 324. O.J.P. Eboli and D. Zeppenfeld, Phys. Lett. **B495**, 147 (2000);  
 H. Davoudiasl, T. Han, and H.E. Logan, Phys. Rev. **D71**, 115007 (2005).  
 325. L. Wang and X.F. Han, Phys. Rev. **D87**, 015015 (2013);  
 K. Schmidt-Hoberg and F. Staub JHEP **1210**, 195 (2012);  
 H. An, T. Liu, and L.T. Wang, Phys. Rev. **D86**, 075030 (2012);  
 D.A. Vasquez *et al.*, Phys. Rev. **D86**, 035023 (2012);  
 S.F. King, M. Muhlleitner, and R. Nevzorov, Nucl. Phys. **B860**, 207 (2012).  
 326. N.D. Christensen, T. Han, Z. Liu and S. Su, JHEP **1308**, 019 (2013);  
 T. Han, Z. Liu and S. Su, JHEP **1408**, 093 (2014);  
 S.F. King *et al.*, Phys. Rev. **D90**, 095014 (2014);  
 N.E. Bomark *et al.*, JHEP **1502**, 044 (2015);  
 P. Athron *et al.*, JHEP **1501**, 153 (2015);  
 N.E. Bomark, S. Moretti and L. Roszkowski, arXiv:1503.04228 [hep-ph] (2015);  
 T. Li and S. Su, JHEP **1511**, 068 (2015);  
 G. Belanger *et al.*, JHEP **1509**, 151 (2015);  
 J. Baglio *et al.*, JHEP **1510**, 024 (2015);  
 M. Guchait and J. Kumar, Int. J. Mod. Phys. **A31**, 1650069 (2016);  
 D. Barducci *et al.*, JHEP **1601**, 050 (2016);  
 M. Carena *et al.*, Phys. Rev. **D93**, 035013 (2016);  
 P. Bandyopadhyay, K. Huitu and S. Niyogi, JHEP **1607**, 015 (2016);  
 U. Ellwanger and M. Rodriguez-Vazquez, JHEP **1602**, 096 (2016);  
 R. Costa *et al.*, JHEP **1606**, 034 (2016);  
 E. Conte *et al.*, JHEP **1605**, 100 (2016).
327. Courtesy of Marcin Barziak based on the public code NMSSMTools,  
[www.th.u-psud.fr/NMDEDECAY/nmssmtools.html](http://www.th.u-psud.fr/NMDEDECAY/nmssmtools.html).  
 328. P. Batra *et al.*, JHEP **0402**, 043 (2004).  
 329. P. Batra *et al.*, JHEP **0406**, 032 (2004);  
 A. Maloney, A. Pierce, and J.G. Wacker, JHEP **0606**, 034 (2006);  
 Y. Zhang *et al.*, Phys. Rev. **D78**, 011302 (2008);  
 C.W. Chiang *et al.*, Phys. Rev. **D81**, 015006 (2010);  
 A.D. Medina, N.R. Shah, and C.E.M. Wagner, Phys. Rev. **D80**, 015001 (2009);  
 M. Endo *et al.*, Phys. Rev. **D85**, 095006 (2012);  
 C. Cheung and H.L. Roberts, JHEP **1312**, 018 (2013).  
 330. R. Huo *et al.*, Phys. Rev. **D87**, 055011 (2013).  
 331. R.T. D'Agnolo, E. Kuflik, and M. Zanetti, JHEP **1303**, 043 (2013).  
 332. A. Azatov and J. Galloway, Int. J. Mod. Phys. **A28**, 1330004 (2013).  
 333. N. Craig and A. Katz, JHEP **1305**, 015 (2013).  
 334. T.F. Feng *et al.*, Nucl. Phys. **B871**, 223 (2013).  
 335. S. Dimopoulos and S. Thomas, Nucl. Phys. **B465**, 23 (1996);  
 S. Thomas, Int. J. Mod. Phys. **A13**, 2307 (1998).  
 336. M.S. Carena *et al.*, Nucl. Phys. **B586**, 92 (2000).  
 337. A. Pilaftsis, Phys. Rev. **D58**, 096010 (1998) and Phys. Lett. **B435**, 88 (1998);  
 K.S. Babu *et al.*, Phys. Rev. **D59**, 016004 (1999).  
 338. G.L. Kane and L.T. Wang, Phys. Lett. **B488**, 383 (2000);  
 S.Y. Choi, M. Drees, and J.S. Lee, Phys. Lett. **B481**, 57 (2000);  
 S.Y. Choi and J.S. Lee, Phys. Rev. **D61**, 015003 (2000);  
 S.Y. Choi, K. Hagiwara, and J.S. Lee, Phys. Rev. **D64**, 032004 (2001) and Phys. Lett. **B529**, 212 (2002);  
 T. Ibrahim and P. Nath, Phys. Rev. **D63**, 035009 (2001);  
 T. Ibrahim, Phys. Rev. **D64**, 035009 (2001);  
 S. Heinemeyer, Eur. Phys. J. **C22**, 521 (2001);  
 S.W. Ham *et al.*, Phys. Rev. **D68**, 055003 (2003).  
 339. M. Frank *et al.*, JHEP **0702**, 047 (2007);  
 S. Heinemeyer *et al.*, Phys. Lett. **B652**, 300 (2007);  
 T. Hahn *et al.*, arXiv:0710.4891 (2007).  
 340. D.A. Demir, Phys. Rev. **D60**, 055006 (1999);  
 S.Y. Choi, M. Drees, and J.S. Lee, Phys. Lett. **B481**, 57 (2000);  
 K. E. Williams, H. Rzehak, and G. Weiglein, Eur. Phys. J. **C71**, 1669 (2011).  
 341. E. Christova *et al.*, Nucl. Phys. **B639**, 263 (2002) [E: Nucl. Phys. **B647**, 359 (2002)].  
 342. M.D. Goodsell and F. Staub, arXiv:1604.05335 [hep-ph] (2016);  
 A. Chakraborty *et al.*, Phys. Rev. **D90**, 055005 (2014);  
 M. Carena *et al.*, JHEP **1602**, 123 (2016).  
 343. J.F. Gunion and H.E. Haber, Phys. Rev. **D67**, 075019 (2003);  
 G.C. Branco *et al.*, Phys. Reports **516**, 1 (2012).  
 344. S.L. Glashow and S. Weinberg, Phys. Rev. **D15**, 1958 (1977);  
 E.A. Paschos, Phys. Rev. **D15**, 1966 (1977);  
 H. Georgi, Hadronic J. **1**, 1227 (1978);  
 H. Haber, G. Kane, and T. Sterling, Nucl. Phys. **B161**, 493 (1979);  
 A.G. Akeroyd, Phys. Lett. **B368**, 89 (1996);  
 A.G. Akeroyd, Nucl. Phys. **B544**, 557 (1999);  
 A.G. Akeroyd, A. Arhrib, and E. Naimi, Eur. Phys. J. **C20**, 51 (2001).  
 345. N.G. Deshpande and E. Ma, Phys. Rev. **D18**, 2574 (1978);  
 R. Barbieri, L.J. Hall, and V. Rychkov, Phys. Rev. **D74**, 015007 (2006);  
 L. Lopez Honorez *et al.*, JCAP **0702**, 028 (2007);  
 E. Lundstrom, M. Gustafsson, and J. Edsjo, Phys. Rev. **D79**, 035013 (2009);  
 E. Dolle *et al.*, Phys. Rev. **D8**, 035003 (2010);  
 X. Miao, S. Su, and B. Thomas, Phys. Rev. **D82**, 035009 (2010);  
 L. Lopez-Honorez and C. Yaguna, JCAP **1101**, 002 (2011).

346. A. Arhrib, R. Benbrik, and N. Gaur, Phys. Rev. **D85**, 095021 (2012);  
B. Swiezewska and M. Krawczyk, Phys. Rev. **D88**, 035019 (2013);  
A. Goudelis, B. Herrmann, and O. Stoel, JHEP **1309**, 106 (2013).
347. V. Barger, H.E. Logan, and G. Shaughnessy, Phys. Rev. **D79**, 115018 (2009).
348. D. O'Connell, M.J. Ramsey-Musolf, and M.B. Wise, Phys. Rev. **D75**, 037701 (2007);  
V. Barger *et al.*, Phys. Rev. **D77**, 035005 (2008);  
V. Barger *et al.*, Phys. Rev. **D79**, 015018 (2009).
349. H.E. Haber, *Proceedings of the 1990 Theoretical Advanced Study Institute in Elementary Particle Physics*, edited by M. Cvetič and Paul Langacker (World Scientific, Singapore, 1991) pp. 340–475 and references therein.
350. S. Glashow and S. Weinberg, Phys. Rev. **D15**, 1958 (1977).
351. G.C. Branco, W. Grimus, and L. Lavoura, Phys. Lett. **B380**, 119 (1996);  
F.J. Botella, G.C. Branco, M.N. Rebelo Phys. Lett. **B687**, 194 (2010).
352. N. Craig and S. Thomas, JHEP **1211**, 083 (2012).
353. G.C. Branco *et al.*, Phys. Reports **516**, 1 (2012).
354. C.Y. Chen, S. Dawson, and M. Sher, Phys. Rev. **D88**, 015018 (2013).
355. B. Swiezewska and M. Krawczyk, Phys. Rev. **D88**, 035019 (2013).
356. B. Coleppa, F. Kling and S. Su, JHEP **1409**, 161 (2014);  
J. Cao *et al.*, JHEP **1412**, 026 (2014);  
B. Coleppa, F. Kling and S. Su, JHEP **1412**, 148 (2014);  
D. Curtin, R. Essig and Y.M. Zhong, JHEP **1506**, 025 (2015);  
F. Kling, A. Pyarelal and S. Su, JHEP **1511**, 051 (2015);  
J. Hajer *et al.*, JHEP **1511**, 124 (2015);  
W. Bernreuther *et al.*, Phys. Rev. **D93**, 034032 (2016);  
S. Gori *et al.*, Phys. Rev. **D93**, 075038 (2016);  
F. Kling, J.M. No and S. Su, arXiv:1604.01406 [hep-ph] (2016);  
N. Craig *et al.*, arXiv:1605.08744 [hep-ph] (2016).
357. N. Craig *et al.*, JHEP **1506**, 137 (2015).
358. J. Schechter and J. W. F. Valle, Phys. Rev. **D22**, 2227 (1980).
359. T.P. Cheng and L.F. Li, Phys. Rev. **D22**, 2860 (1980).
360. H. Georgi and M. Machacek, Nucl. Phys. **B262**, 463 (1985).
361. M.S. Chanowitz and M. Golden, Phys. Lett. **B165**, 105 (1985).
362. H.E. Logan and M.A. Roy, Phys. Rev. **D82**, 115011 (2010).
363. H.E. Haber and H.E. Logan, Phys. Rev. **D62**, 015011 (2000).
364. A.G. Akeroyd, M. Aoki, and H. Sugiyama, Phys. Rev. **D77**, 075010 (2008).
365. P. Nath *et al.*, Nucl. Phys. (Proc. Supp.) **B200**, 185 (2010).
366. J. Garayoa and T. Schwetz, JHEP **0803**, 009 (2008).
367. J.F. Gunion, R. Vega, and J. Wudka, Phys. Rev. **D43**, 2322 (1991).
368. S. Kanemura and K. Yagyu, Phys. Rev. **D85**, 115009 (2012).
369. I. Low and J. Lykken, JHEP **1010**, 053 (2010).
370. C. Englert, E. Re, and M. Spannowsky, Phys. Rev. **D87**, 095014 (2013).
371. I. Low, J. Lykken, and G. Shaughnessy, Phys. Rev. **D86**, 093012 (2012).
372. A. Falkowski, S. Rychkov, and A. Urbano, JHEP **1204**, 073 (2012).
373. B.A. Dobrescu and J.D. Lykken, JHEP **1302**, 073 (2013).
374. D. Carmi *et al.*, JHEP **1210**, 196 (2012).
375. C.W. Chiang and K. Yagyu, JHEP **1301**, 026 (2013).
376. G. Belanger *et al.*, Phys. Rev. **D88**, 075008 (2013).
377. C. Englert, E. Re, and M. Spannowsky, Phys. Rev. **D88**, 035024 (2013).
378. B. Bellazzini, C. Csaki and J. Serra, Eur. Phys. J. **C74**, 2766 (2014).
379. N. Arkani-Hamed *et al.*, JHEP **0207**, 034 (2002).
380. N. Arkani-Hamed, A.G. Cohen, and H. Georgi, Phys. Lett. **B513**, 232 (2001).
381. N. Arkani-Hamed *et al.*, JHEP **0208**, 021 (2002).
382. M. Schmaltz, JHEP **0408**, 056 (2004).
383. M. Schmaltz, D. Stolarski, and J. Thaler, JHEP **1009**, 018 (2010).
384. M. Perelstein, Prog. in Part. Nucl. Phys. **58**, 247 (2007).
385. M. Schmaltz and D. Tucker-Smith, Ann. Rev. Nucl. and Part. Sci. **55**, 229 (2005).
386. J.A. Casas, J.R. Espinosa, and I. Hidalgo, JHEP **0503**, 038 (2005).
387. H.C. Cheng and I. Low, JHEP **0309**, 051 (2003).
388. M.S. Carena *et al.*, Phys. Rev. **D75**, 091701 (2007).
389. ATLAS Collab., ATLAS-CONF-2012-147, ATLAS-CONF-2012-109 and ATLAS-CONF-2013-024 (2012);  
CMS Collab., CMS-PAS-EXO-12-048 and CMS-SUS-12-028 (2012).
390. CMS Collab., CMS-PAS-B2G-12-015 (2012);  
ATLAS Collab., ATLAS-CONF-2013-060 (2013).
391. R.S. Chivukula, M. Narain, and J. Womersley, *Dynamical Electroweak Symmetry Breaking*, in this volume.
392. H. Georgi, A.E. Nelson, and A. Manohar, Phys. Lett. **B126**, 169 (1983);  
A.E. Nelson and M.J. Strassler, JHEP **0009**, 030 (2000);  
S. Davidson, G. Isidori, and S. Uhlig, Phys. Lett. **B663**, 73 (2008).
393. C. Csaki, A. Falkowski, and A. Weiler, JHEP **0809**, 008 (2008);  
B. Keren-Zur *et al.*, Nucl. Phys. **B867**, 429 (2013).
394. K. Agashe, R. Contino, and A. Pomarol, Nucl. Phys. **B719**, 165 (2005).
395. O. Matsedonskyi, G. Panico, and A. Wulzer, JHEP **1301**, 164 (2013);  
M. Redi and A. Tesi, JHEP **1210**, 166 (2012);  
D. Marzocca, M. Serone, and J. Shu, JHEP **1208**, 013 (2012);  
A. Pomarol and F. Riva, JHEP **1208**, 135 (2012).
396. R. Contino and G. Servant, JHEP **0806**, 026 (2008);  
J. Mrazek and A. Wulzer, Phys. Rev. **D81**, 075006 (2010).
397. A. De Simone *et al.*, JHEP **1304**, 004 (2013);  
A. Azatov *et al.*, Phys. Rev. **D89**, 075001 (2014).
398. A. Falkowski, Phys. Rev. **D77**, 055018 (2008);  
I. Low and A. Vichi, Phys. Rev. **D84**, 045019 (2011);  
A. Azatov and J. Galloway, Phys. Rev. **D85**, 055013 (2012);  
C. Delaunay, C. Grojean, and G. Perez, JHEP **1309**, 090 (2013).
399. R.K. Ellis *et al.*, Nucl. Phys. **B297**, 221 (1988);  
U. Baur and E.W.N. Glover, Nucl. Phys. **B339**, 38 (1990);  
O. Brein and W. Hollik, Phys. Rev. **D68**, 095006 (2003);  
U. Langenegger *et al.*, JHEP **0606**, 035 (2006).
400. A. Banfi, A. Martin, and V. Sanz, JHEP **1408**, 053 (2014);  
A. Azatov and A. Paul, JHEP **1401**, 014 (2014);  
C. Grojean *et al.*, JHEP **1405**, 022 (2014);  
M. Schlaffer *et al.*, Eur. Phys. J. **C74**, 3120 (2014);  
M. Buschmann *et al.*, Phys. Rev. **D90**, 013010 (2014);  
M. Buschmann *et al.*, JHEP **1502**, 038 (2015);  
U. Langenegger, M. Spira and I. Strebel, arXiv:1507.01373 [hep-ph] (2015).
401. A. Azatov *et al.*, J. Exp. Theor. Phys. **120**, 354 (2015).
402. A. Azatov *et al.*, Phys. Rev. **D92**, 035001 (2015).
403. R. Contino, L. Da Rold, and A. Pomarol, Phys. Rev. **D75**, 055014 (2007).
404. D. Pappadopulo, A. Thamm, and R. Torre, JHEP **1307**, 058 (2013);  
M. Montull *et al.*, Phys. Rev. **D88**, 095006 (2013).
405. I. Low, R. Rattazzi, and A. Vichi, JHEP **1004**, 126 (2010).
406. M. Ciuchini *et al.*, JHEP **1308**, 106 (2013).
407. C. Grojean, O. Matsedonskyi, and G. Panico, JHEP **1310**, 160 (2013).
408. M. Geller and O. Telem, Phys. Rev. Lett. **114**, 191801 (2015).
409. P. Batra and Z. Chacko, Phys. Rev. **D79**, 095012 (2009);  
R. Barbieri *et al.*, JHEP **1508**, 161 (2015);  
M. Low, A. Tesi and L. T. Wang, Phys. Rev. **D91**, 095012 (2015).
410. B. Bellazzini *et al.*, Eur. Phys. J. **C74**, 2790 (2014);  
F. Coradeschi *et al.*, JHEP **1311**, 057 (2013).

411. CMS Collab., *Submitted to Phys. Lett. B*, CERN-PH-EP-2015-211 (2015).
412. ATLAS Collab., *Phys. Rev.* **D92**, 092004 (2015).
413. CMS Collab., CMS-PAS-HIG-13-032 (2013).
414. ATLAS Collab., ATLAS-CONF-2013-027 (2013).
415. CMS Collab., CMS-PAS-HIG-13-016 (2013).
416. ATLAS Collab., ATLAS-CONF-2012-013 (2012).
417. CMS Collab., CMS-PAS-HIG-12-002 (2012).
418. CMS Collab., CMS-HIG-12-013 (2013), CERN-PH-EP/2013-011.
419. ALEPH Collab., *Phys. Lett.* **B526**, 191 (2002).
420. L3 Collab., *Phys. Lett.* **B545**, 30 (2002).
421. S. Schael *et al.*, [ALEPH, DELPHI, L3 and OPAL Collaborations and LEP Working Group for Higgs Boson Searches], *Eur. Phys. J.* **C47**, 547 (2006).
422. M.M. Kado and C.G. Tully, *Ann. Rev. Nucl. and Part. Sci.* **52**, 65 (2002).
423. OPAL Collab., *Eur. Phys. J.* **C23**, 397 (2002).
424. DELPHI Collab., *Eur. Phys. J.* **C38**, 1 (2004).
425. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Lett.* **B698**, 97 (2011).
426. T. Aaltonen *et al.*, [CDF Collab.], *Phys. Rev.* **D85**, 032005 (2012).
427. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Rev. Lett.* **104**, 151801 (2010).
428. D0 Collab., D0 Note 5974-CONF (2011).
429. ATLAS Collab., ATLAS-CONF-2012-094 (2012).
430. CMS Collab., CMS-PAS-HIG-13-021 (2013).
431. S. Chatrchyan *et al.*, [CMS Collab.], *Phys. Lett.* **B722**, 207 (2013).
432. CMS Collab., CMS-PAS-HIG-16-007 (2016).
433. CMS Collab., *Submitted to Phys. Lett. B*, CERN-PH-EP/2015-284 (2015).
434. J. Bernon *et al.*, *Phys. Rev.* **D91**, 075019 (2015).
435. ATLAS Collab., *JHEP* **08**, 148 (2015).
436. S. Chatrchyan *et al.*, [CMS Collab.], *JHEP* **09**, 029 (2012).
437. A. Djouadi, J. Ellis and J. Quevillon, [arXiv:1605.00542](https://arxiv.org/abs/1605.00542) [hep-ph] (2016);  
M Carena and Zhen Liu, to appear.
438. ATLAS Collab., *Phys. Lett.* **B744**, 163 (2015).
439. S.H. Zhu, [hep-ph/9901221](https://arxiv.org/abs/hep-ph/9901221) (1999).
440. H.E. Logan and S. Su, *Phys. Rev.* **D66**, 035001 (2002).
441. A. Gutierrez-Rodriguez and O.A. Sampayo, *Phys. Rev.* **D62**, 055004 (2000).
442. S. Kanemura, S. Moretti, and K. Odagiri, *JHEP* **0102**, 011 (2001).
443. B. Abbott *et al.*, [D0 Collab.], *Phys. Rev. Lett.* **82**, 4975 (1999).
444. A. Abulencia *et al.*, [CDF Collab.], *Phys. Rev. Lett.* **96**, 042003 (2006).
445. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Lett.* **B682**, 278 (2009).
446. ATLAS Collab., *JHEP* **1206**, 039 (2012).
447. ATLAS Collab., *JHEP* **03**, 088 (2015).
448. CMS Collab., CMS-PAS-HIG-11-019.
449. CMS Collab., *Submitted to JHEP*, CERN-PH-EP-2015-221 (2015).
450. ATLAS Collab., *Submitted to JHEP*, CERN-PH-EP-2015-290 (2015).
451. ATLAS Collab., ATLAS-CONF-2012-010 (2012).
452. CMS Collab., *Submitted to JHEP*, CERN-PH-EP-2015-266 (2015).
453. ATLAS Collab., *Phys. Rev. Lett.* **114**, 231801 (2015).
454. J.F. Gunion, R. Vega, and J. Wudka, *Phys. Rev.* **D42**, 1673 (1990).
455. J. Abdallah *et al.*, [DELPHI Collab.], *Eur. Phys. J.* **C54**, 1 (2008) [E: **C56**, 165 (2008)].
456. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Rev. Lett.* **103**, 061801 (2009).
457. R. Dermisek, *Mod. Phys. Lett.* **A24**, 1631 (2009).
458. J.F. Gunion, *JHEP* **0908**, 032 (2009).
459. U. Ellwanger, *Eur. Phys. J.* **C71**, 1782 (2011).
460. W. Love *et al.*, [CLEO Collab.], *Phys. Rev. Lett.* **101**, 151802 (2008).
461. B. Aubert *et al.*, [BaBar Collab.], *Phys. Rev. Lett.* **103**, 081803 (2009).
462. B. Aubert *et al.*, [BaBar Collab.], *Phys. Rev. Lett.* **103**, 181801 (2009).
463. V.M. Abazov *et al.*, [D0 Collab.], *Phys. Rev. Lett.* **103**, 061801 (2009).
464. ATLAS Collab., ATLAS-CONF-2011-020 (2011).
465. S. Chatrchyan *et al.*, [CMS Collab.], *Phys. Rev. Lett.* **109**, 121801 (2012).
466. ATLAS Collab., ATLAS-CONF-2012-079 (2012).
467. CMS Collab., *Phys. Lett.* **B752**, 146 (2016).
468. CMS Collab., *JHEP* **01**, 079 (2016).
469. S. Schael *et al.*, [ALEPH Collab.], *JHEP* **1005**, 049 (2010).
470. ATLAS Collab., ATLAS-CONF-2013-013 (2013).
471. S. Chatrchyan *et al.*, [CMS Collab.], *Eur. Phys. J.* **C73**, 2469 (2013).
472. ATLAS Collab., ATLAS-CONF-2012-016 (2012).
473. CMS Collab., CMS-PAS-HIG-13-014 (2013).
474. ATLAS Collab., ATLAS-CONF-2012-017 (2012).
475. ATLAS Collab., ATLAS-CONF-2012-163 (2012).
476. CMS Collab., CMS-PAS-HIG-12-024 (2012).
477. ATLAS Collab., ATLAS-CONF-2013-067 (2013).
478. ATLAS Collab., ATLAS-CONF-2012-018 (2012).
479. CMS Collab., CMS-PAS-HIG-13-008 (2013).
480. CMS Collab., CMS-PAS-HIG-12-046 (2012).
481. R. Aaij *et al.*, [LHCb Collab.], *JHEP* **1305**, 132 (2013).
482. ATLAS Collab., *Phys. Rev.* **D92**, 052002 (2015).
483. G. Aad *et al.*, [ATLAS Collab.], *Eur. Phys. J.* **C72**, 2244 (2012).
484. S. Chatrchyan *et al.*, [CMS Collab.], *Eur. Phys. J.* **C72**, 2189 (2012).
485. ATLAS Collab., CERN-PH-EP-2013-172 (2013).
486. ATLAS Collab., ATLAS-CONF-2013-081 (2013).
487. M. Carena, C. Grojean, M. Kado, and V. Sharma, *Status of Higgs boson physics*, in *Review of Particle Physics*, *Chin. Phys.* **C38** (2014) 090001.

## 12. THE CKM QUARK-MIXING MATRIX

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### 12.1. Introduction

The masses and mixings of quarks have a common origin in the Standard Model (SM). They arise from the Yukawa interactions with the Higgs condensate,

$$\mathcal{L}_Y = -Y_{ij}^d \overline{Q}_{Li}^d \phi d_{Rj}^I - Y_{ij}^u \overline{Q}_{Li}^u \epsilon \phi^* u_{Rj}^I + \text{h.c.}, \quad (12.1)$$

where  $Y^{u,d}$  are  $3 \times 3$  complex matrices,  $\phi$  is the Higgs field,  $i, j$  are generation labels, and  $\epsilon$  is the  $2 \times 2$  antisymmetric tensor.  $Q_L^I$  are left-handed quark doublets, and  $d_R^I$  and  $u_R^I$  are right-handed down- and up-type quark singlets, respectively, in the weak-eigenstate basis. When  $\phi$  acquires a vacuum expectation value,  $\langle \phi \rangle = (0, v/\sqrt{2})$ , Eq. (12.1) yields mass terms for the quarks. The physical states are obtained by diagonalizing  $Y^{u,d}$  by four unitary matrices,  $V_{L,R}^{u,d}$ , as  $M_{\text{diag}}^f = V_L^f Y^f V_R^{f\dagger} (v/\sqrt{2})$ ,  $f = u, d$ . As a result, the charged-current  $W^\pm$  interactions couple to the physical  $u_{Lj}$  and  $d_{Lk}$  quarks with couplings given by

$$\frac{-g}{\sqrt{2}} (\overline{u}_L, \overline{c}_L, \overline{t}_L) \gamma^\mu W_\mu^+ V_{\text{CKM}} \begin{pmatrix} d_L \\ s_L \\ b_L \end{pmatrix} + \text{h.c.},$$

$$V_{\text{CKM}} \equiv V_L^u V_L^{d\dagger} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix}. \quad (12.2)$$

This Cabibbo-Kobayashi-Maskawa (CKM) matrix [1,2] is a  $3 \times 3$  unitary matrix. It can be parameterized by three mixing angles and the  $CP$ -violating KM phase [2]. Of the many possible conventions, a standard choice has become [3]

$$V_{\text{CKM}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$= \begin{pmatrix} c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{-i\delta} \\ -s_{12} c_{23} - c_{12} s_{23} s_{13} e^{i\delta} & c_{12} c_{23} - s_{12} s_{23} s_{13} e^{i\delta} & s_{23} c_{13} \\ s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta} & c_{23} c_{13} \end{pmatrix}, \quad (12.3)$$

where  $s_{ij} = \sin \theta_{ij}$ ,  $c_{ij} = \cos \theta_{ij}$ , and  $\delta$  is the phase responsible for all  $CP$ -violating phenomena in flavor-changing processes in the SM. The angles  $\theta_{ij}$  can be chosen to lie in the first quadrant, so  $s_{ij}, c_{ij} \geq 0$ .

It is known experimentally that  $s_{13} \ll s_{23} \ll s_{12} \ll 1$ , and it is convenient to exhibit this hierarchy using the Wolfenstein parameterization. We define [4–6]

$$s_{12} = \lambda = \frac{|V_{us}|}{\sqrt{|V_{ud}|^2 + |V_{us}|^2}}, \quad s_{23} = A\lambda^2 = \lambda \left| \frac{V_{cb}}{V_{us}} \right|,$$

$$s_{13} e^{i\delta} = V_{ub}^* = A\lambda^3 (\bar{\rho} + i\bar{\eta}) = \frac{A\lambda^3 (\bar{\rho} + i\bar{\eta}) \sqrt{1 - A^2 \lambda^4}}{\sqrt{1 - \lambda^2} [1 - A^2 \lambda^4 (\bar{\rho} + i\bar{\eta})]}. \quad (12.4)$$

These relations ensure that  $\bar{\rho} + i\bar{\eta} = -(V_{ud} V_{ub}^*) / (V_{cd} V_{cb}^*)$  is phase convention independent, and the CKM matrix written in terms of  $\lambda$ ,  $A$ ,  $\bar{\rho}$ , and  $\bar{\eta}$  is unitary to all orders in  $\lambda$ . The definitions of  $\bar{\rho}, \bar{\eta}$  reproduce all approximate results in the literature. For example,  $\bar{\rho} = \rho(1 - \lambda^2/2 + \dots)$  and one can write  $V_{\text{CKM}}$  to  $\mathcal{O}(\lambda^4)$  either in terms of  $\bar{\rho}, \bar{\eta}$  or, traditionally,

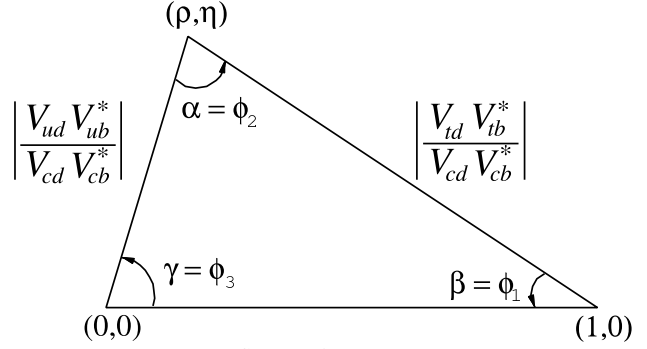


Figure 12.1: Sketch of the unitarity triangle.

$$V_{\text{CKM}} = \begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\bar{\rho} - i\bar{\eta}) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \bar{\rho} - i\bar{\eta}) & -A\lambda^2 & 1 \end{pmatrix} + \mathcal{O}(\lambda^4). \quad (12.5)$$

The CKM matrix elements are fundamental parameters of the SM, so their precise determination is important. The unitarity of the CKM matrix imposes  $\sum_i V_{ij} V_{ik}^* = \delta_{jk}$  and  $\sum_j V_{ij} V_{kj}^* = \delta_{ik}$ . The six vanishing combinations can be represented as triangles in a complex plane, of which those obtained by taking scalar products of neighboring rows or columns are nearly degenerate. The areas of all triangles are the same, half of the Jarlskog invariant,  $J$  [7], which is a phase-convention-independent measure of  $CP$  violation, defined by  $\text{Im}[V_{ij} V_{kl} V_{il}^* V_{kj}^*] = J \sum_{m,n} \epsilon_{ikm} \epsilon_{jln}$ .

The most commonly used unitarity triangle arises from

$$V_{ud} V_{ub}^* + V_{cd} V_{cb}^* + V_{td} V_{tb}^* = 0, \quad (12.6)$$

by dividing each side by the best-known one,  $V_{cd} V_{cb}^*$  (see Fig. 1). Its vertices are exactly  $(0,0)$ ,  $(1,0)$ , and, due to the definition in Eq. (12.4),  $(\bar{\rho}, \bar{\eta})$ . An important goal of flavor physics is to overconstrain the CKM elements, and many measurements can be conveniently displayed and compared in the  $\bar{\rho}, \bar{\eta}$  plane. While the Lagrangian in Eq. (12.1) is renormalized, and the CKM matrix has a well known scale dependence above the weak scale [8], below  $\mu = m_W$  the CKM elements can be treated as constants, with all  $\mu$ -dependence contained in the running of quark masses and higher-dimension operators.

Unless explicitly stated otherwise, we describe all measurements assuming the SM, to extract magnitudes and phases of CKM elements in Sec. 12.2 and 12.3. Processes dominated by loop-level contributions in the SM are particularly sensitive to new physics. We give the global fit results for the CKM elements in Sec. 12.4, and discuss some implications for beyond standard model physics in Sec. 12.5.

### 12.2. Magnitudes of CKM elements

#### 12.2.1. $|V_{ud}|$ :

The most precise determination of  $|V_{ud}|$  comes from the study of superallowed  $0^+ \rightarrow 0^+$  nuclear beta decays, which are pure vector transitions. Taking the average of the fourteen most precise determinations [9] yields

$$|V_{ud}| = 0.97417 \pm 0.00021. \quad (12.7)$$

The error is dominated by theoretical uncertainties stemming from nuclear Coulomb distortions and radiative corrections. A precise determination of  $|V_{ud}|$  is also obtained from the measurement of the neutron lifetime. The theoretical uncertainties are very small, but the determination is limited by the knowledge of the ratio of the axial-vector and vector couplings,  $g_A = G_A/G_V$  [10]. The PIBETA experiment [11] has improved the measurement of the  $\pi^+ \rightarrow \pi^0 e^+ \nu$  branching ratio to 0.6%, and quotes  $|V_{ud}| = 0.9728 \pm 0.0030$ , in agreement with the more precise result listed above. The interest in this measurement is that the determination of  $|V_{ud}|$  is very clean theoretically, because it is a pure vector transition and is free from nuclear-structure uncertainties.



### 12.2.2. $|V_{us}|$ :

The product of  $|V_{us}|$  and the form factor at  $q^2 = 0$ ,  $|V_{us}|f_+(0)$ , has been extracted traditionally from  $K_L^0 \rightarrow \pi e \nu$  decays in order to avoid isospin-breaking corrections ( $\pi^0 - \eta$  mixing) that affect  $K^\pm$  semileptonic decay, and the complications induced by a second (scalar) form factor present in the muonic decays. The last round of measurements has lead to enough experimental constraints to justify the comparison between different decay modes. Systematic errors related to the experimental quantities, *e.g.*, the lifetime of neutral or charged kaons, and the form factor determinations for electron and muonic decays, differ among decay modes, and the consistency between different determinations enhances the confidence in the final result. For this reason, we follow the prescription [12] to average  $K_L^0 \rightarrow \pi e \nu$ ,  $K_L^0 \rightarrow \pi \mu \nu$ ,  $K^\pm \rightarrow \pi^0 e^\pm \nu$ ,  $K^\pm \rightarrow \pi^0 \mu^\pm \nu$  and  $K_S^0 \rightarrow \pi e \nu$ . The average of these five decay modes yields  $|V_{us}|f_+(0) = 0.2165 \pm 0.0004$ . Results obtained from each decay mode, and exhaustive references to the experimental data, are listed for instance in Ref. [10]. The form factor average  $f_+(0) = 0.9677 \pm 0.0037$  [13] from three-flavor lattice QCD calculations gives  $|V_{us}| = 0.2237 \pm 0.0009$  [10].<sup>1</sup> The broadly used classic calculation of  $f_+(0)$  [16] is in good agreement with this value, while other calculations [17] differ by as much as 2%.

The calculation of the ratio of the kaon and pion decay constants enables one to extract  $|V_{us}/V_{ud}|$  from  $K \rightarrow \mu \nu(\gamma)$  and  $\pi \rightarrow \mu \nu(\gamma)$ , where  $(\gamma)$  indicates that radiative decays are included [18]. The KLOE measurement of the  $K \rightarrow \mu \nu(\gamma)$  branching ratio [19], combined with the lattice QCD result,  $f_K/f_\pi = 1.1928 \pm 0.0026$  [13], leads to  $|V_{us}| = 0.2254 \pm 0.0008$ , where the accuracy is limited by the knowledge of the ratio of the decay constants. The average of these two determinations is quoted as [10]

$$|V_{us}| = 0.2248 \pm 0.0006. \quad (12.8)$$

The latest determination from hyperon decays can be found in Ref. [20]. The authors focus on the analysis of the vector form factor, protected from first order  $SU(3)$  breaking effects by the Ademollo-Gatto theorem [21], and treat the ratio between the axial and vector form factors  $g_1/f_1$  as experimental input, thus avoiding first order  $SU(3)$  breaking effects in the axial-vector contribution. They find  $|V_{us}| = 0.2250 \pm 0.0027$ , although this does not include an estimate of the theoretical uncertainty due to second-order  $SU(3)$  breaking, contrary to Eq. (12.8). Concerning hadronic  $\tau$  decays to strange particles, averaging both inclusive and exclusive  $\tau \rightarrow h \nu$  ( $h = \pi, K$ ) decay measurements yield  $|V_{us}| = 0.2204 \pm 0.0014$  [22,23].

### 12.2.3. $|V_{cd}|$ :

The magnitude of  $V_{cd}$  can be extracted from semileptonic charm decays, using theoretical knowledge of the form factors. In semileptonic  $D$  decays, lattice QCD calculations have predicted the normalization of the  $D \rightarrow \pi \ell \nu$  and  $D \rightarrow K \ell \nu$  form factors [14]. The dependence on the invariant mass of the lepton pair,  $q^2$ , is determined from lattice QCD and theoretical constraints from analyticity [15]. Using three-flavor lattice QCD calculations for  $D \rightarrow \pi \ell \nu$ ,  $f_+^{D\pi}(0) = 0.666 \pm 0.029$  [14], and the average [22,24] of measurements of recent BaBar [25] and BESIII [26] as well as CLEO-c [27] and Belle [28] of  $D \rightarrow \pi \ell \nu$  decays, one obtains  $|V_{cd}| = 0.214 \pm 0.003 \pm 0.009$ , where the first uncertainty is experimental, and the second is from the theoretical uncertainty of the form factor.

The determination of  $|V_{cd}|$  is also possible from leptonic decay  $D^+ \rightarrow \mu^+ \nu$ . Its precision has been improved by a recent BESIII measurement [29]. Averaged with earlier CLEO measurement [30] and  $f_D = 209.2 \pm 3.3$  MeV [14], one obtains  $|V_{cd}| = 0.219 \pm 0.005 \pm 0.003$ .

Earlier determinations of  $|V_{cd}|$  came from neutrino scattering data. The difference of the ratio of double-muon to single-muon production by neutrino and antineutrino beams is proportional to the charm cross section off valence  $d$  quarks, and therefore to  $|V_{cd}|^2$  times the average

semileptonic branching ratio of charm mesons,  $\mathcal{B}_\mu$ . The method was used first by CDHS [31] and then by CCFR [32,33] and CHARM II [34]. Averaging these results is complicated, because it requires assumptions about the scale of the QCD corrections, and because  $\mathcal{B}_\mu$  is an effective quantity, which depends on the specific neutrino beam characteristics. Given that no recent experimental input is available, we quote the average from a past review,  $\mathcal{B}_\mu |V_{cd}|^2 = (0.463 \pm 0.034) \times 10^{-2}$  [35]. Analysis cuts make these experiments insensitive to neutrino energies smaller than 30 GeV. Thus,  $\mathcal{B}_\mu$  should be computed using only neutrino interactions with visible energy larger than 30 GeV. An appraisal [36] based on charm-production fractions measured in neutrino interactions [37,38] gives  $\mathcal{B}_\mu = 0.088 \pm 0.006$ . Data from the CHORUS experiment [39] are sufficiently precise to extract  $\mathcal{B}_\mu$  directly, by comparing the number of charm decays with a muon to the total number of charmed hadrons found in the nuclear emulsions. Requiring the visible energy to be larger than 30 GeV, CHORUS finds  $\mathcal{B}_\mu = 0.085 \pm 0.009 \pm 0.006$ . We use the average of these two determinations,  $\mathcal{B}_\mu = 0.087 \pm 0.005$ , and obtain  $|V_{cd}| = 0.230 \pm 0.011$ . Averaging the three determinations above, we find

$$|V_{cd}| = 0.220 \pm 0.005. \quad (12.9)$$

### 12.2.4. $|V_{cs}|$ :

The direct determination of  $|V_{cs}|$  is possible from semileptonic  $D$  or leptonic  $D_s$  decays, using lattice QCD calculations of the semileptonic  $D$  form factor or the  $D_s$  decay constant. For muonic decays, the average of Belle [40], CLEO-c [41] and BABAR [42] is  $\mathcal{B}(D_s^+ \rightarrow \mu^+ \nu) = (5.56 \pm 0.24) \times 10^{-3}$  [43]. For decays to  $\tau$  leptons, the average of CLEO-c [41,44,45], BABAR [42] and Belle [40] gives  $\mathcal{B}(D_s^+ \rightarrow \tau^+ \nu) = (5.56 \pm 0.22) \times 10^{-2}$  [43]. From each of these values, determinations of  $|V_{cs}|$  can be obtained using the PDG values for the mass and lifetime of the  $D_s$ , the masses of the leptons, and  $f_{D_s} = (248.6 \pm 2.7)$  MeV [14]. The average of these determinations gives  $|V_{cs}| = 1.008 \pm 0.021$ , where the error is dominated by the lattice QCD determination of  $f_{D_s}$ . In semileptonic  $D$  decays, lattice QCD calculations of the  $D \rightarrow K \ell \nu$  form factor are available [14]. Using  $f_+^{DK}(0) = 0.747 \pm 0.019$  and the average of CLEO-c [27], Belle [28], BABAR [46] and recent BESIII [26] measurements of  $D \rightarrow K \ell \nu$  decays, one obtains  $|V_{cs}| = 0.975 \pm 0.007 \pm 0.025$ , where the first error is experimental and the second, which is dominant, is from the theoretical uncertainty of the form factor. Averaging the determinations from leptonic and semileptonic decays, we find

$$|V_{cs}| = 0.995 \pm 0.016. \quad (12.10)$$

Measurements of on-shell  $W^\pm$  decays sensitive to  $|V_{cs}|$  were made by LEP-2. The  $W$  branching ratios depend on the six CKM elements involving quarks lighter than  $m_W$ . The  $W$  branching ratio to each lepton flavor is  $1/\mathcal{B}(W \rightarrow \ell \bar{\nu}_\ell) = 3[1 + \sum_{u,c,d,s,b} |V_{ij}|^2 (1 + \alpha_s(m_W)/\pi) + \dots]$ . Assuming lepton universality, the measurement  $\mathcal{B}(W \rightarrow \ell \bar{\nu}_\ell) = (10.83 \pm 0.07 \pm 0.07)\%$  [47] implies  $\sum_{u,c,d,s,b} |V_{ij}|^2 = 2.002 \pm 0.027$ . This is a precise test of unitarity; however, only flavor-tagged  $W$ -decays determine  $|V_{cs}|$  directly, such as DELPHI's tagged  $W^+ \rightarrow c \bar{s}$  analysis, yielding  $|V_{cs}| = 0.94_{-0.26}^{+0.32} \pm 0.13$  [48].

### 12.2.5. $|V_{cb}|$ :

This matrix element can be determined from exclusive and inclusive semileptonic decays of  $B$  mesons to charm. The inclusive determinations use the semileptonic decay rate measurement, together with (certain moments of) the leptonic energy and the hadronic invariant-mass spectra. The theoretical basis is the operator product expansion [49,50], which allows calculation of the decay rate and various spectra as expansions in  $\alpha_s$  and inverse powers of the heavy-quark mass. The dependence on  $m_b$ ,  $m_c$ , and the parameters that occur at subleading order is different for different moments, and a large number of measured moments overconstrains all the parameters, and tests the consistency of the determination. The precise extraction of  $|V_{cb}|$  requires using a “threshold” quark mass definition [51,52]. Inclusive measurements have been performed using  $B$  mesons from  $Z^0$  decays at LEP, and at  $e^+e^-$  machines operated at the  $\Upsilon(4S)$ . At LEP, the large boost of  $B$  mesons from the  $Z^0$  allows the determination of

<sup>1</sup> For lattice QCD inputs, we use the averages from Ref. [14] whenever possible, unless the minireviews [10,15] choose other values. We only use unquenched lattice QCD results. Hereafter, the first error is statistical and the second is systematic, unless mentioned otherwise.

the moments throughout phase space, which is not possible otherwise, but the large statistics available at the  $B$  factories lead to more precise determinations. An average of the measurements and a compilation of the references are provided by Ref. [15]:  $|V_{cb}| = (42.2 \pm 0.8) \times 10^{-3}$ .

Exclusive determinations are based on semileptonic  $B$  decays to  $D$  and  $D^*$ . In the  $m_{b,c} \gg \Lambda_{\text{QCD}}$  limit, all form factors are given by a single Isgur-Wise function [53], which depends on the product of the four-velocities of the  $B$  and  $D^{(*)}$  mesons,  $w = v \cdot v'$ . Heavy-quark symmetry determines the rate at  $w = 1$ , the maximum momentum transfer to the leptons, and  $|V_{cb}|$  is obtained from an extrapolation to  $w = 1$ . The exclusive determination,  $|V_{cb}| = (39.2 \pm 0.7) \times 10^{-3}$  [15], has a comparable precision to the inclusive one, and the main theoretical uncertainty in the form factor and the experimental uncertainty in the rate near  $w = 1$  are to a large extent independent of the inclusive determination. The  $V_{cb}$  and  $V_{ub}$  minireview [15] quotes a combination with the error scaled by  $\sqrt{\chi^2} = 2.9$ ,

$$|V_{cb}| = (40.5 \pm 1.5) \times 10^{-3}. \quad (12.11)$$

Less precise measurements of  $|V_{cb}|$ , not included in this average, can be obtained from  $\mathcal{B}(B \rightarrow D^{(*)}\tau\bar{\nu})$ . The most precise data involving  $\tau$  modes are the  $|V_{cb}|$ -independent ratios,  $\mathcal{B}(B \rightarrow D^{(*)}\tau\bar{\nu})/\mathcal{B}(B \rightarrow D^{(*)}\ell\bar{\nu})$  [54]. If the currently nearly  $4\sigma$  hint of lepton non-universality is confirmed, the determination of  $|V_{cb}|$  becomes more complicated.

### 12.2.6. $|V_{ub}|$ :

The determination of  $|V_{ub}|$  from inclusive  $B \rightarrow X_u \ell \bar{\nu}$  decay is complicated due to large  $B \rightarrow X_c \ell \bar{\nu}$  backgrounds. In most regions of phase space where the charm background is kinematically forbidden, the hadronic physics enters via unknown nonperturbative functions, so-called shape functions. (In contrast, the nonperturbative physics for  $|V_{cb}|$  is encoded in a few parameters.) At leading order in  $\Lambda_{\text{QCD}}/m_b$ , there is only one shape function, which can be extracted from the photon energy spectrum in  $B \rightarrow X_s \gamma$  [55,56], and applied to several spectra in  $B \rightarrow X_u \ell \bar{\nu}$ . The subleading shape functions are modeled in the current determinations. Phase space cuts for which the rate has only subleading dependence on the shape function are also possible [57]. The measurements of both the hadronic and the leptonic systems are important for an optimal choice of phase space. A different approach is to make the measurements more inclusive by extending them deeper into the  $B \rightarrow X_c \ell \bar{\nu}$  region, and thus reduce the theoretical uncertainties. Analyses of the electron-energy endpoint from CLEO [58], BABAR [59], and Belle [60] quote  $B \rightarrow X_u e \bar{\nu}$  partial rates for  $|\vec{p}_e| \geq 2.0 \text{ GeV}$  and  $1.9 \text{ GeV}$ , which are well below the charm endpoint. The large and pure  $B\bar{B}$  samples at the  $B$  factories permit the selection of  $B \rightarrow X_u \ell \bar{\nu}$  decays in events where the other  $B$  is fully reconstructed [61]. With this full-reconstruction tag method, the four-momenta of both the leptonic and the hadronic final states can be measured. It also gives access to a wider kinematic region, because of improved signal purity. Ref. [15] quotes the inclusive average,  $|V_{ub}| = (4.49 \pm 0.16_{-0.18}^{+0.16}) \times 10^{-3}$ .

To extract  $|V_{ub}|$  from exclusive decays, the form factors have to be known. Experimentally, better signal-to-background ratios are offset by smaller yields. The  $B \rightarrow \pi \ell \bar{\nu}$  branching ratio is now known to 5%. Lattice QCD calculations of the  $B \rightarrow \pi \ell \bar{\nu}$  form factor are available [62,63] for the high  $q^2$  region ( $q^2 > 16$  or  $18 \text{ GeV}^2$ ). A fit to the experimental partial rates and lattice results versus  $q^2$  yields  $|V_{ub}| = (3.72 \pm 0.16) \times 10^{-3}$  [63]. Light-cone QCD sum rules are supposed to be applicable for  $q^2 < 12 \text{ GeV}^2$  [64]. The minireview [15] quotes a combination,  $|V_{ub}| = (3.72 \pm 0.19) \times 10^{-3}$ .

The uncertainties in extracting  $|V_{ub}|$  from inclusive and exclusive decays are different to a large extent. A combination of the determinations is quoted [15] with the error scaled by  $\sqrt{\chi^2} = 2.6$ ,

$$|V_{ub}| = (4.09 \pm 0.39) \times 10^{-3}. \quad (12.12)$$

A determination of  $|V_{ub}|$  not included in this average can be obtained from  $\mathcal{B}(B \rightarrow \tau \bar{\nu}) = (1.06 \pm 0.20) \times 10^{-4}$  [43]. Using  $f_B = (190.5 \pm 4.2) \text{ MeV}$  [14] and  $\tau_{B^\pm} = (1.638 \pm 0.004) \text{ ps}$  [65], we find  $|V_{ub}| = (4.04 \pm 0.38) \times 10^{-3}$ . This decay is sensitive, for example,

to tree-level charged Higgs contributions, and the measured rate is consistent with the SM expectation. The recent LHCb measurement  $|V_{ub}/V_{cb}| = 0.083 \pm 0.006$  [66] from the ratio of  $\Lambda_b \rightarrow p^+ \mu^- \bar{\nu}$  and  $\Lambda_b \rightarrow \Lambda_c^+ \mu^- \bar{\nu}$  in different regions of  $q^2$ , will hopefully be averaged with the above, using more than one lattice QCD inputs, by the next edition.

### 12.2.7. $|V_{td}|$ and $|V_{ts}|$ :

The CKM elements  $|V_{td}|$  and  $|V_{ts}|$  are not likely to be precisely measurable in tree-level processes involving top quarks, so one has to rely on determinations from  $B-\bar{B}$  oscillations mediated by box diagrams with top quarks, or loop-mediated rare  $K$  and  $B$  decays. Theoretical uncertainties in hadronic effects limit the accuracy of the current determinations. These can be reduced by taking ratios of processes that are equal in the flavor  $SU(3)$  limit to determine  $|V_{td}/V_{ts}|$ .

The mixing of the two  $B^0$  mesons was discovered by ARGUS [67], and the mass difference is precisely measured by now,  $\Delta m_d = (0.5064 \pm 0.0019) \text{ ps}^{-1}$  [68]. In the  $B_s^0$  system,  $\Delta m_s$  was first measured significantly by CDF [69] and the world average, dominated by a recent LHCb measurement [70], is  $\Delta m_s = (17.757 \pm 0.021) \text{ ps}^{-1}$  [68]. Neglecting corrections suppressed by  $|V_{tb}| - 1$ , and using the lattice QCD results  $f_{B_d} \sqrt{\hat{B}_{B_d}} = (216 \pm 15) \text{ MeV}$  and  $f_{B_s} \sqrt{\hat{B}_{B_s}} = (266 \pm 18) \text{ MeV}$  [14],

$$|V_{td}| = (8.2 \pm 0.6) \times 10^{-3}, \quad |V_{ts}| = (40.0 \pm 2.7) \times 10^{-3}. \quad (12.13)$$

The uncertainties are dominated by lattice QCD. Several uncertainties are reduced in the calculation of the ratio  $\xi = (f_{B_s} \sqrt{\hat{B}_{B_s}})/(f_{B_d} \sqrt{\hat{B}_{B_d}}) = 1.268 \pm 0.063$  [14] and therefore the constraint on  $|V_{td}/V_{ts}|$  from  $\Delta m_d/\Delta m_s$  is more reliable theoretically. These provide a theoretically clean and significantly improved constraint

$$|V_{td}/V_{ts}| = 0.215 \pm 0.001 \pm 0.011. \quad (12.14)$$

The inclusive branching ratio  $\mathcal{B}(B \rightarrow X_s \gamma) = (3.43 \pm 0.22) \times 10^{-4}$  extrapolated to  $E_\gamma > E_0 = 1.6 \text{ GeV}$  [71] is also sensitive to  $|V_{tb}V_{ts}|$ . In addition to  $t$ -quark penguins, a substantial part of the rate comes from charm contributions proportional to  $V_{cb}V_{cs}^*$  via the application of  $3 \times 3$  CKM unitarity (which is used here). With the NNLO calculation of  $\mathcal{B}(B \rightarrow X_s \gamma)_{E_\gamma > E_0}/\mathcal{B}(B \rightarrow X_c e \bar{\nu})$  [72], we obtain  $|V_{ts}/V_{cb}| = 0.99 \pm 0.05$ . The  $B_s \rightarrow \mu^+ \mu^-$  rate is also proportional to  $|V_{tb}V_{ts}|^2$  in the SM, and the observed signal  $\mathcal{B}(B_s \rightarrow \mu^+ \mu^-) = (2.8_{-0.6}^{+0.7}) \times 10^{-9}$  [73] is consistent with the SM, with sizable uncertainties.

A complementary determination of  $|V_{td}/V_{ts}|$  is possible from the ratio of  $B \rightarrow \rho \gamma$  and  $K^* \gamma$  rates. The ratio of the neutral modes is theoretically cleaner than that of the charged ones, because the poorly known spectator-interaction contribution is expected to be smaller ( $W$ -exchange vs. weak annihilation). For now, because of low statistics, we average the charged and neutral rates assuming the isospin symmetry and heavy-quark limit motivated relation,  $|V_{td}/V_{ts}|^2/\xi_\gamma^2 = [\Gamma(B^+ \rightarrow \rho^+ \gamma) + 2\Gamma(B^0 \rightarrow \rho^0 \gamma)]/[\Gamma(B^+ \rightarrow K^{*+} \gamma) + \Gamma(B^0 \rightarrow K^{*0} \gamma)] = (3.19 \pm 0.46)\%$  [71]. Here  $\xi_\gamma$  contains the poorly known hadronic physics. Using  $\xi_\gamma = 1.2 \pm 0.2$  [74], and combining the experimental and theoretical errors in quadrature, gives  $|V_{td}/V_{ts}| = 0.214 \pm 0.016 \pm 0.036$ .

A theoretically clean determination of  $|V_{td}V_{ts}^*|$  is possible from  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  decay [75]. Experimentally, only seven events have been observed [76] and the rate is consistent with the SM with large uncertainties. Much more data are needed for a precision measurement.

### 12.2.8. $|V_{tb}|$ :

The determination of  $|V_{tb}|$  from top decays uses the ratio of branching fractions  $R = \mathcal{B}(t \rightarrow Wb)/\mathcal{B}(t \rightarrow Wq) = |V_{tb}|^2/(\sum_q |V_{tq}|^2) = |V_{tb}|^2$ , where  $q = b, s, d$ . The CDF and DØ measurements performed on data collected during Run II of the Tevatron give  $|V_{tb}| > 0.78$  [77] and  $0.99 > |V_{tb}| > 0.90$  [78], respectively, at 95% CL. CMS measured the same quantity at 7 TeV and gives  $|V_{tb}| > 0.92$  [79] at 95% CL.

The direct determination of  $|V_{tb}|$ , without assuming unitarity, is possible from the single top-quark-production cross section. The  $(3.30^{+0.52}_{-0.40})$  pb combined cross section [80] of  $D\bar{O}$  and CDF measurements implies  $|V_{tb}| = 1.02^{+0.06}_{-0.05}$ . The LHC experiments, ATLAS and CMS, have measured single-top production cross sections (and extracted  $|V_{tb}|$ ) in  $t$ -channel,  $Wt$ -channel, and  $s$ -channel at 7 TeV, 8 TeV, and 13 TeV [81]. The average of these  $|V_{tb}|$  values is calculated to be  $|V_{tb}| = 1.005 \pm 0.036$ , where all systematic errors and theoretical errors are treated to be fully correlated. The average of Tevatron and LHC values gives

$$|V_{tb}| = 1.009 \pm 0.031. \quad (12.15)$$

The experimental systematic uncertainties dominate, and a dedicated combination would be welcome.

A weak constraint on  $|V_{tb}|$  can be obtained from precision electroweak data, where top quarks enter in loops. The sensitivity is best in  $\Gamma(Z \rightarrow b\bar{b})$  and yields  $|V_{tb}| = 0.77^{+0.18}_{-0.24}$  [82].

## 12.3. Phases of CKM elements

As can be seen from Fig. 12.1, the angles of the unitarity triangle are

$$\begin{aligned} \beta = \phi_1 &= \arg\left(-\frac{V_{cd}V_{cb}^*}{V_{td}V_{tb}^*}\right), \\ \alpha = \phi_2 &= \arg\left(-\frac{V_{td}V_{tb}^*}{V_{ud}V_{ub}^*}\right), \\ \gamma = \phi_3 &= \arg\left(-\frac{V_{ud}V_{ub}^*}{V_{cd}V_{cb}^*}\right). \end{aligned} \quad (12.16)$$

Since  $CP$  violation involves phases of CKM elements, many measurements of  $CP$ -violating observables can be used to constrain these angles and the  $\bar{\rho}, \bar{\eta}$  parameters.

### 12.3.1. $\epsilon$ and $\epsilon'$

The measurement of  $CP$  violation in  $K^0-\bar{K}^0$  mixing,  $|\epsilon| = (2.233 \pm 0.015) \times 10^{-3}$  [83], provides important information about the CKM matrix. The phase of  $\epsilon$  is determined by long-distance physics,  $\epsilon = \frac{1}{2} e^{i\phi_\epsilon} \sin \phi_\epsilon \arg(-M_{12}/\Gamma_{12})$ , where  $\phi_\epsilon = \arctan |2\Delta m_K/\Delta\Gamma_K| \simeq 43.5^\circ$ . The SM prediction can be written as

$$\begin{aligned} \epsilon &= \kappa_\epsilon e^{i\phi_\epsilon} \frac{G_F^2 m_W^2 m_K}{12\sqrt{2}\pi^2 \Delta m_K} f_K^2 \hat{B}_K \left\{ \eta_{tt} S(x_t) \text{Im}[(V_{ts}V_{td}^*)^2] \right. \\ &\quad \left. + 2\eta_{ct} S(x_c, x_t) \text{Im}(V_{cs}V_{cd}^* V_{ts}V_{td}^*) + \eta_{cc} x_c \text{Im}[(V_{cs}V_{cd}^*)^2] \right\}, \end{aligned} \quad (12.17)$$

where  $\kappa_\epsilon \simeq 0.94 \pm 0.02$  [84] includes the effects of  $\Delta s = 1$  operators and  $\phi_\epsilon \neq \pi/4$  (see also Ref. [85]). The displayed terms are the short-distance  $\Delta s = 2$  contribution to  $\text{Im}M_{12}$  in the usual phase convention,  $S$  is an Inami-Lim function [86],  $x_q = m_q^2/m_W^2$ , and  $\eta_{ij}$  are perturbative QCD corrections. The constraint from  $\epsilon$  in the  $\bar{\rho}, \bar{\eta}$  plane is bounded by approximate hyperbolas. Lattice QCD determined the bag parameter  $\hat{B}_K = 0.766 \pm 0.010$  [14], and the main uncertainties now come from  $(V_{ts}V_{td}^*)^2$ , which is approximately  $\sigma(|V_{cb}|^4) \sim \sigma(A^4)$ , the  $\eta_{ij}$  coefficients, and estimates of  $\kappa_\epsilon$ .

The measurement of  $6\text{Re}(\epsilon'/\epsilon) = 1 - |\eta_{00}/\eta_{+-}|^2$ , where each  $\eta_{ij} = \langle \pi^i \pi^j | \mathcal{H} | K_L \rangle / \langle \pi^i \pi^j | \mathcal{H} | K_S \rangle$  violates  $CP$ , provides a qualitative test of the CKM mechanism, and strong constraints on many new physics scenarios. Its nonzero value,  $\text{Re}(\epsilon'/\epsilon) = (1.67 \pm 0.23) \times 10^{-3}$  [83], demonstrated the existence of direct  $CP$  violation, a prediction of the KM ansatz. While  $\text{Re}(\epsilon'/\epsilon) \propto \text{Im}(V_{td}V_{ts}^*)$ , this quantity cannot easily be used to extract CKM parameters, because the electromagnetic penguin contributions tend to cancel the gluonic penguins for large  $m_t$  [87], thus enhancing hadronic uncertainties. Most SM estimates [88–91] agree with the observed value, indicating that  $\bar{\eta}$  is positive. Progress in lattice QCD [92] may eventually yield a precise SM prediction.

### 12.3.2. $\beta / \phi_1$

#### 12.3.2.1. Charmonium modes:

$CP$ -violation measurements in  $B$ -meson decays provide direct information on the angles of the unitarity triangle, shown in Fig. 12.1. These overconstraining measurements serve to improve the determination of the CKM elements, or to reveal effects beyond the SM.

The time-dependent  $CP$  asymmetry of neutral  $B$  decays to a final state  $f$  common to  $B^0$  and  $\bar{B}^0$  is given by [93,94]

$$\mathcal{A}_f = \frac{\Gamma(\bar{B}^0(t) \rightarrow f) - \Gamma(B^0(t) \rightarrow f)}{\Gamma(\bar{B}^0(t) \rightarrow f) + \Gamma(B^0(t) \rightarrow f)} = S_f \sin(\Delta m_d t) - C_f \cos(\Delta m_d t), \quad (12.18)$$

where

$$S_f = \frac{2\text{Im}\lambda_f}{1 + |\lambda_f|^2}, \quad C_f = \frac{1 - |\lambda_f|^2}{1 + |\lambda_f|^2}, \quad \lambda_f = \frac{q}{p} \frac{\bar{A}_f}{A_f}. \quad (12.19)$$

Here,  $q/p$  describes  $B^0-\bar{B}^0$  mixing and, to a good approximation in the SM,  $q/p = V_{tb}^*V_{td}/V_{tb}V_{td}^* = e^{-2i\beta + \mathcal{O}(\lambda^4)}$  in the usual phase convention.  $A_f$  ( $\bar{A}_f$ ) is the amplitude of the  $B^0 \rightarrow f$  ( $\bar{B}^0 \rightarrow f$ ) decay. If  $f$  is a  $CP$  eigenstate, and amplitudes with one CKM phase dominate the decay, then  $|A_f| = |\bar{A}_f|$ ,  $C_f = 0$ , and  $S_f = \sin(\arg \lambda_f) = \eta_f \sin 2\phi$ , where  $\eta_f$  is the  $CP$  eigenvalue of  $f$  and  $2\phi$  is the phase difference between the  $B^0 \rightarrow f$  and  $B^0 \rightarrow \bar{B}^0 \rightarrow f$  decay paths. A contribution of another amplitude to the decay with a different CKM phase makes the value of  $S_f$  sensitive to relative strong-interaction phases between the decay amplitudes (it also makes  $C_f \neq 0$  possible).

The  $b \rightarrow c\bar{c}s$  decays to  $CP$  eigenstates ( $B^0 \rightarrow \text{charmonium } K_{S,L}^0$ ) are the theoretically cleanest examples, measuring  $S_f = -\eta_f \sin 2\beta$ . The  $b \rightarrow s q \bar{q}$  penguin amplitudes have dominantly the same weak phase as the  $b \rightarrow c\bar{c}s$  tree amplitude. Since only  $\lambda^2$ -suppressed penguin amplitudes introduce a new  $CP$ -violating phase, amplitudes with a single weak phase dominate, and we expect  $|\bar{A}_{\psi K}/A_{\psi K}| - 1| < 0.01$ . The  $e^+e^-$  asymmetric-energy  $B$ -factory experiments, *BABAR* [95] and *Belle* [96], provide precise measurements. The world average including LHCb [97] and other measurements is [98]

$$\sin 2\beta = 0.691 \pm 0.017. \quad (12.20)$$

This measurement has a four-fold ambiguity in  $\beta$ , which can be resolved by a global fit as mentioned in Sec. 12.4. Experimentally, the two-fold ambiguity  $\beta \rightarrow \pi/2 - \beta$  (but not  $\beta \rightarrow \pi + \beta$ ) can be resolved by a time-dependent angular analysis of  $B^0 \rightarrow J/\psi K^{*0}$  [99,100], or a time-dependent Dalitz plot analysis of  $B^0 \rightarrow \bar{D}^0 h^0$  ( $h^0 = \pi^0, \eta, \omega$ ) with  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  [101,102]. These results indicate that negative  $\cos 2\beta$  solutions are very unlikely, in agreement with the global CKM fit result.

The  $b \rightarrow c\bar{c}d$  mediated transitions, such as  $B^0 \rightarrow J/\psi \pi^0$  and  $B^0 \rightarrow D^{(*)+} D^{(*)-}$ , also measure approximately  $\sin 2\beta$ . However, the dominant component of the  $b \rightarrow d$  penguin amplitude has a different CKM phase ( $V_{tb}^*V_{td}$ ) than the tree amplitude ( $V_{cb}^*V_{cd}$ ), and its magnitudes are of the same order in  $\lambda$ . Therefore, the effect of penguins could be large, resulting in  $S_f \neq -\eta_f \sin 2\beta$  and  $C_f \neq 0$ . These decay modes have also been measured by *BABAR* and *Belle*. The world averages [98],  $S_{J/\psi \pi^0} = -0.93 \pm 0.15$ ,  $S_{J/\psi \rho^0} = -0.66^{+0.16}_{-0.12}$ ,  $S_{D^+ D^-} = -0.98 \pm 0.17$ , and  $S_{D^{*+} D^{*-}} = -0.71 \pm 0.09$  ( $\eta_f = +1$  for these modes), are consistent with  $\sin 2\beta$  obtained from  $B^0 \rightarrow \text{charmonium } K^0$  decays, and the  $C_f$ 's are consistent with zero, although the uncertainties are sizable.

The  $b \rightarrow c\bar{u}d$  decays,  $B^0 \rightarrow \bar{D}^0 h^0$  with  $\bar{D}^0 \rightarrow CP$  eigenstates, have no penguin contributions and provide theoretically clean  $\sin 2\beta$  measurements. The joint analysis of *BABAR* and *Belle* gives  $S_{D^{(*)} h^0} = -0.66 \pm 0.12$  [103].

### 12.3.3.2. Penguin-dominated modes:

The  $b \rightarrow \bar{s}q$  penguin-dominated decays have the same CKM phase as the  $b \rightarrow c\bar{s}$  tree level decays, up to corrections suppressed by  $\lambda^2$ , since  $V_{tb}^*V_{ts} = -V_{cb}^*V_{cs}[1 + \mathcal{O}(\lambda^2)]$ . Therefore, decays such as  $B^0 \rightarrow \phi K^0$  and  $\eta' K^0$  provide  $\sin 2\beta$  measurements in the SM. Any new physics contribution to the amplitude with a different weak phase would give rise to  $S_f \neq -\eta_f \sin 2\beta$ , and possibly  $C_f \neq 0$ . Therefore, the main interest in these modes is not simply to measure  $\sin 2\beta$ , but to search for new physics. Measurements of many other decay modes in this category, such as  $B \rightarrow \pi^0 K_S^0$ ,  $K_S^0 K_S^0 K_S^0$ , etc., have also been performed by BABAR and Belle. The results and their uncertainties are summarized in Fig. 12.3 and Table 12.1 of Ref. [94].

### 12.3.3. $\alpha / \phi_2$ :

Since  $\alpha$  is the phase between  $V_{tb}^*V_{td}$  and  $V_{ub}^*V_{ud}$ , only time-dependent  $CP$  asymmetries in  $b \rightarrow u\bar{u}d$  decay dominated modes can directly measure  $\sin 2\alpha$ , in contrast to  $\sin 2\beta$ , where several different transitions can be used. Since  $b \rightarrow d$  penguin amplitudes have a different CKM phase than  $b \rightarrow u\bar{u}d$  tree amplitudes, and their magnitudes are of the same order in  $\lambda$ , the penguin contribution can be sizable, which makes the determination of  $\alpha$  complicated. To date,  $\alpha$  has been measured in  $B \rightarrow \pi\pi$ ,  $\rho\pi$  and  $\rho\rho$  decay modes.

#### 12.3.3.1. $B \rightarrow \pi\pi$ :

It is now experimentally well established that there is a sizable contribution of  $b \rightarrow d$  penguin amplitudes in  $B \rightarrow \pi\pi$  decays. Thus,  $S_{\pi^+\pi^-}$  in the time-dependent  $B^0 \rightarrow \pi^+\pi^-$  analysis does not measure  $\sin 2\alpha$ , but

$$S_{\pi^+\pi^-} = \sqrt{1 - C_{\pi^+\pi^-}^2} \sin(2\alpha + 2\Delta\alpha), \quad (12.21)$$

where  $2\Delta\alpha$  is the phase difference between  $e^{2i\gamma}\bar{A}_{\pi^+\pi^-}$  and  $A_{\pi^+\pi^-}$ . The value of  $\Delta\alpha$ , hence  $\alpha$ , can be extracted using the isospin relation among the amplitudes of  $B^0 \rightarrow \pi^+\pi^-$ ,  $B^0 \rightarrow \pi^0\pi^0$ , and  $B^+ \rightarrow \pi^+\pi^0$  decays [104],

$$\frac{1}{\sqrt{2}} A_{\pi^+\pi^-} + A_{\pi^0\pi^0} - A_{\pi^+\pi^0} = 0, \quad (12.22)$$

and a similar expression for the  $\bar{A}_{\pi\pi}$ 's. This method utilizes the fact that a pair of pions from  $B \rightarrow \pi\pi$  decay must be in a zero angular momentum state, and, because of Bose statistics, they must have even isospin. Consequently,  $\pi^0\pi^\pm$  is in a pure isospin-2 state, while the penguin amplitudes only contribute to the isospin-0 final state. The latter does not hold for the electroweak penguin amplitudes, but their effect is expected to be small. The isospin analysis uses the world averages of BABAR, Belle and LHCb measurements [98]  $S_{\pi^+\pi^-} = -0.66 \pm 0.06$ ,  $C_{\pi^+\pi^-} = -0.31 \pm 0.05$ , the branching fractions of all three modes, and the direct  $CP$  asymmetry  $C_{\pi^0\pi^0} = -0.43^{+0.25}_{-0.24}$ . This analysis leads to 16 mirror solutions for  $0 \leq \alpha < 2\pi$ . Because of this, and the sizable experimental error of the  $B^0 \rightarrow \pi^0\pi^0$  rate and  $CP$  asymmetry, only a loose constraint on  $\alpha$  can be obtained at present [105],  $0^\circ < \alpha < 3.8^\circ$ ,  $86.2^\circ < \alpha < 102.9^\circ$ ,  $122.1^\circ < \alpha < 147.9^\circ$ , and  $167.1^\circ < \alpha < 180^\circ$  at 68% CL.

#### 12.3.3.2. $B \rightarrow \rho\rho$ :

The decay  $B^0 \rightarrow \rho^+\rho^-$  contains two vector mesons in the final state, which in general is a mixture of  $CP$ -even and  $CP$ -odd components. Therefore, it was thought that extracting  $\alpha$  from this mode would be complicated.

However, the longitudinal polarization fractions ( $f_L$ ) in  $B^+ \rightarrow \rho^+\rho^0$  and  $B^0 \rightarrow \rho^+\rho^-$  decays were measured to be close to unity [106], which implies that the final states are almost purely  $CP$ -even. Furthermore,  $\mathcal{B}(B^0 \rightarrow \rho^0\rho^0) = (0.97 \pm 0.24) \times 10^{-6}$  is much smaller than  $\mathcal{B}(B^0 \rightarrow \rho^+\rho^-) = (24.2^{+3.1}_{-3.2}) \times 10^{-6}$  and  $\mathcal{B}(B^+ \rightarrow \rho^+\rho^0) = (24.0^{+1.9}_{-2.0}) \times 10^{-6}$  [22], which implies that the effect of the penguin diagrams is small. The isospin analysis using the world averages,  $S_{\rho^+\rho^-} = -0.14 \pm 0.13$  and  $C_{\rho^+\rho^-} = -0.00 \pm 0.09$  [22], together with the time-dependent  $CP$  asymmetry,  $S_{\rho^0\rho^0} = -0.3 \pm 0.7$  and  $C_{\rho^0\rho^0} = -0.2 \pm 0.9$  [107], and the above mentioned branching fractions, gives  $0^\circ < \alpha < 5.6^\circ$ ,  $84.4^\circ < \alpha < 95.3^\circ$  and  $174.7^\circ < \alpha < 180^\circ$  at 68% CL [105], with mirror solutions at  $3\pi/2 - \alpha$ . A possible small violation of Eq. (12.22) due to the finite width of the  $\rho$  [108] is neglected.

### 12.3.3.3. $B \rightarrow \rho\pi$ :

The final state in  $B^0 \rightarrow \rho^+\pi^-$  decay is not a  $CP$  eigenstate, but this decay proceeds via the same quark-level diagrams as  $B^0 \rightarrow \pi^+\pi^-$ , and both  $B^0$  and  $\bar{B}^0$  can decay to  $\rho^+\pi^-$ . Consequently, mixing-induced  $CP$  violations can occur in four decay amplitudes,  $B^0 \rightarrow \rho^+\pi^-$  and  $\bar{B}^0 \rightarrow \rho^\pm\pi^\mp$ . The time-dependent Dalitz plot analysis of  $B^0 \rightarrow \pi^+\pi^-\pi^0$  decays permits the extraction of  $\alpha$  with a single discrete ambiguity,  $\alpha \rightarrow \alpha + \pi$ , since one knows the variation of the strong phases in the interference regions of the  $\rho^+\pi^-$ ,  $\rho^-\pi^+$ , and  $\rho^0\pi^0$  amplitudes in the Dalitz plot [109]. The combination of Belle [110] and BABAR [111] measurements gives  $\alpha = (54.1^{+7.7}_{-10.3})^\circ$  and  $(141.8^{+4.7}_{-5.4})^\circ$  [105]. This constraint is still moderate.

Combining the  $B \rightarrow \pi\pi$ ,  $\rho\pi$ , and  $\rho\rho$  decay modes [105],  $\alpha$  is constrained as

$$\alpha = (87.6^{+3.5}_{-3.3})^\circ. \quad (12.23)$$

A different statistical approach [112] gives similar constraint from the combination of these measurements.

### 12.3.4. $\gamma / \phi_3$ :

By virtue of Eq. (12.16),  $\gamma$  does not depend on CKM elements involving the top quark, so it can be measured in tree-level  $B$  decays. This is an important distinction from the measurements of  $\alpha$  and  $\beta$ , and implies that the measurements of  $\gamma$  are unlikely to be affected by physics beyond the SM.

#### 12.3.4.1. $B^\pm \rightarrow DK^\pm$ :

The interference of  $B^- \rightarrow D^0 K^-$  ( $b \rightarrow c\bar{u}s$ ) and  $B^- \rightarrow \bar{D}^0 K^-$  ( $b \rightarrow u\bar{c}s$ ) transitions can be studied in final states accessible in both  $D^0$  and  $\bar{D}^0$  decays [93]. In principle, it is possible to extract the  $B$  and  $D$  decay amplitudes, the relative strong phases, and the weak phase  $\gamma$  from the data.

A practical complication is that the precision depends sensitively on the ratio of the interfering amplitudes

$$r_B = \left| A(B^- \rightarrow \bar{D}^0 K^-) / A(B^- \rightarrow D^0 K^-) \right|, \quad (12.24)$$

which is around 0.1–0.2. The original GLW method [113,114] considers  $D$  decays to  $CP$  eigenstates, such as  $B^\pm \rightarrow D_{CP}^{(*)}(\rightarrow \pi^+\pi^-)K^\pm$ . To alleviate the smallness of  $r_B$  and make the interfering amplitudes (which are products of the  $B$  and  $D$  decay amplitudes) comparable in magnitude, the ADS method [115] considers final states where Cabibbo-allowed  $\bar{D}^0$  and doubly-Cabibbo-suppressed  $D^0$  decays interfere. Extensive measurements [98] have been made by the  $B$  factories, CDF and LHCb using both methods.

It was realized that both  $D^0$  and  $\bar{D}^0$  have large branching fractions to certain three-body final states, such as  $K_S\pi^+\pi^-$ , and the analysis can be optimized by studying the Dalitz plot dependence of the interferences [116,117]. The best present determination of  $\gamma$  comes from this method. Belle [118] and BABAR [119] obtained  $\gamma = (78^{+11}_{-12} \pm 4 \pm 9)^\circ$  and  $\gamma = (68 \pm 14 \pm 4 \pm 3)^\circ$ , respectively, where the last uncertainty is due to the  $D$ -decay modeling. LHCb also measured  $\gamma = (62^{+15}_{-14})^\circ$  with the Dalitz model independent manner [120]. The error is sensitive to the central value of the amplitude ratio  $r_B$  (and  $r_B^*$  for the  $D^*K$  mode), for which Belle found somewhat larger central values than BABAR and LHCb. The same values of  $r_B^{(*)}$  enter the ADS analyses, and the data can be combined to fit for  $r_B^{(*)}$  and  $\gamma$ . The  $D^0$ – $\bar{D}^0$  mixing has been neglected in all measurements, but its effect on  $\gamma$  is far below the present experimental accuracy [121], unless  $D^0$ – $\bar{D}^0$  mixing is due to  $CP$ -violating new physics, in which case it can be included in the analysis [122].

Combining the GLW, ADS, and Dalitz analyses [105],  $\gamma$  is constrained as

$$\gamma = (73.2^{+6.3}_{-7.0})^\circ. \quad (12.25)$$

Similar results are found in Ref. [112].



$A_{\text{SL}}^{d,s}$ , are important [138]. A  $D\bar{O}$  measurement sensitive to certain linear combinations of  $A_{\text{SL}}^d$  and  $A_{\text{SL}}^s$  shows a  $3.6\sigma$  hint of a deviation from the SM [139].

Many key measurements which are sensitive to BSM flavor physics are not useful to think about in terms of constraining the unitarity triangle in Fig. 12.1. For example, besides the angles in Eq. (12.16), a key quantity in the  $B_s$  system is  $\beta_s = \arg(-V_{ts}V_{tb}^*/V_{cs}V_{cb}^*)$ , which is the small,  $\lambda^2$ -suppressed, angle of a “squashed” unitarity triangle, obtained by taking the scalar product of the second and third columns. This angle can be measured via time-dependent  $CP$  violation in  $B_s^0 \rightarrow J/\psi\phi$ , similar to  $\beta$  in  $B^0 \rightarrow J/\psi K^0$ . Since the  $J/\psi\phi$  final state is not a  $CP$  eigenstate, an angular analysis of the decay products is needed to separate the  $CP$ -even and  $CP$ -odd components, which give opposite asymmetries. In the SM, the asymmetry for the  $CP$ -even part is  $2\beta_s$  (sometimes the notation  $\phi_s = -2\beta_s$  plus a possible BSM contribution to the  $B_s$  mixing phase is used). Testing if the data agree with the SM prediction,  $2\beta_s = 0.0363 \pm 0.0018$  [105], is another sensitive test of the SM. After the first Tevatron  $CP$ -asymmetry measurements of  $B_s^0 \rightarrow J/\psi\phi$  hinted at a possible tension with the SM, the current world average, dominated by LHCb [140] including  $B_s \rightarrow J/\psi K^+ K^-$  and  $J/\psi \pi^+ \pi^-$  measurements, is  $2\beta_s = 0.034 \pm 0.033$  [22]. This uncertainty is about 20 times the SM uncertainty; thus a lot will be learned from higher-precision measurements in the future.

In the kaon sector, the two measured  $CP$ -violating observables  $\epsilon$  and  $\epsilon'$  are tiny, so models in which all sources of  $CP$  violation are small were viable before the  $B$ -factory measurements. Since the measurement of  $\sin 2\beta$ , we know that  $CP$  violation can be an  $\mathcal{O}(1)$  effect, and only flavor mixing is suppressed between the three quark generations. Thus, many models with spontaneous  $CP$  violation are excluded. In the kaon sector, a very clean test of the SM will come from measurements of  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  and  $K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$ . These loop-induced rare decays are sensitive to new physics, and will allow a determination of  $\beta$ , independent of its value measured in  $B$  decays [141].

The CKM elements are fundamental parameters, so they should be measured as precisely as possible. The overconstraining measurements of  $CP$  asymmetries, mixing, semileptonic, and rare decays severely constrain the magnitudes and phases of possible new physics contributions to flavor-changing interactions. If new particles are observed at the LHC, it will be important to explore their flavor parameters as precisely as possible to understand the underlying physics.

#### References:

- N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
- M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
- L.L. Chau and W.Y. Keung, Phys. Rev. Lett. **53**, 1802 (1984).
- L. Wolfenstein, Phys. Rev. Lett. **51**, 1945 (1983).
- A.J. Buras *et al.*, Phys. Rev. **D50**, 3433 (1994) [hep-ph/9403384].
- J. Charles *et al.* [CKMfitter Group], Eur. Phys. J. **C41**, 1 (2005) [hep-ph/0406184].
- C. Jarlskog, Phys. Rev. Lett. **55**, 1039 (1985).
- W.J. Marciano and A. Sirlin, Nucl. Phys. **B93**, 303 (1975); K.S. Babu, Z. Phys. **C35**, 69 (1987).
- J.C. Hardy and I. S. Towner, Phys. Rev. **C91**, 025501 (2015) [arXiv:1411.5987].
- E. Blucher and W.J. Marciano, “ $V_{ud}$ ,  $V_{us}$ , the Cabibbo Angle and CKM Unitarity,” in this Review.
- D. Poganic *et al.*, Phys. Rev. Lett. **93**, 181803 (2004) [hep-ex/0312030].
- M. Antonelli *et al.* [The FlaviaNet Kaon Working Group], arXiv:0801.1817; see also <http://www.lnf.infn.it/wg/vus>.
- J.L. Rosner, S. Stone, R.S. Van de Water, “Leptonic Decays of Charged Pseudoscalar Mesons,” in this Review.
- S. Aoki *et al.*, “Review of lattice results concerning low energy particle physics,” Eur. Phys. J. **C74**, 2890 (2014) arXiv:1310.8555, <http://itpwiki.unibe.ch/flag>.
- R. Kowalewski and T. Mannel, “Determination of  $V_{cb}$  and  $V_{ub}$ ,” in this Review.
- H. Leutwyler and M. Roos, Z. Phys. **C25**, 91 (1984).
- J. Bijnens and P. Talavera, Nucl. Phys. **B669**, 341 (2003) [hep-ph/0303103]; M. Jamin *et al.*, JHEP **402**, 047 (2004) [hep-ph/0401080]; V. Cirigliano *et al.*, JHEP **504**, 6 (2005) [hep-ph/0503108]; C. Dawson *et al.*, PoS **LAT2005**, 337 (2005) [hep-lat/0510018]; N. Tsutsui *et al.* [JLQCD Collab.], PoS **LAT2005**, 357 (2005) [hep-lat/0510068]; M. Okamoto [Fermilab Lattice Collab.], hep-lat/0412044.
- W.J. Marciano, Phys. Rev. Lett. **93**, 231803 (2004) [hep-ph/0402299].
- F. Ambrosino *et al.* [KLOE Collab.], Phys. Lett. **B632**, 76 (2006) [hep-ex/0509045].
- N. Cabibbo *et al.*, Ann. Rev. Nucl. and Part. Sci. **53**, 39 (2003) [hep-ph/0307298]; Phys. Rev. Lett. **92**, 251803 (2004) [hep-ph/0307214].
- M. Ademollo and R. Gatto, Phys. Rev. Lett. **13**, 264 (1964).
- Y. Amhis *et al.*, Heavy Flavor Averaging Group, arXiv:1412.7515, and updates at <http://www.slac.stanford.edu/xorg/hfag/>.
- Heavy Flavor Averaging Group [22] and Tau Summer 2014 report [ $V_{us}$ ] measurement, <http://www.slac.stanford.edu/xorg/hfag/tau/summer-2014/vus.html>.
- Heavy Flavor Averaging Group [22], Charm Semileptonic Decays, October 2014, [http://www.slac.stanford.edu/xorg/hfag/charm/semilep2014/charm\\_semilep.htm](http://www.slac.stanford.edu/xorg/hfag/charm/semilep2014/charm_semilep.htm).
- J. P. Lees *et al.* [BABAR Collab.], Phys. Rev. **D91**, 052022 (2015) [arXiv:1412.5502].
- Y.H. Zhang, presented at ICHEP 2014 [for BES III Collab.].
- D. Besson *et al.* [CLEO Collab.], Phys. Rev. **D80**, 032005 (2009) [arXiv:0906.2983].
- L. Widhalm *et al.* [Belle Collab.], Phys. Rev. Lett. **97**, 061804 (2006) [hep-ex/0604049].
- M. Ablikim *et al.* [BESIII Collab.], Phys. Rev. **D89**, 051104 (2014) [arXiv:1312.0374].
- B. I. Eisenstein *et al.* [CLEO Collab.], Phys. Rev. **D78**, 052003 (2008) [arXiv:0806.2112].
- H. Abramowicz *et al.* [CHDS Collab.], Z. Phys. **C15**, 19 (1982).
- S.A. Rabinowitz *et al.* [CCFR Collab.], Phys. Rev. Lett. **70**, 134 (1993).
- A.O. Bazarko *et al.* [CCFR Collab.], Z. Phys. **C65**, 189 (1995) [hep-ex/9406007].
- P. Vilain *et al.* [CHARM II Collab.], Eur. Phys. J. **C11**, 19 (1999).
- F.J. Gilman *et al.*, Phys. Lett. **B592**, 793 (2004).
- G.D. Lellis *et al.*, Phys. Rept. **399**, 227 (2004) [Erratum *ibid.* **411**, 323 (2005)].
- N. Ushida *et al.* [Fermilab E531 Collab.], Phys. Lett. **B206**, 380 (1988).
- T. Bolton, hep-ex/9708014.
- A. Kayis-Topaksu *et al.* [CHORUS Collab.], Phys. Lett. **B626**, 24 (2005).
- A. Zupanc *et al.* [Belle Collab.], JHEP **1309**, 139 (2013) [arXiv:1307.6240].
- J.P. Alexander *et al.* [CLEO Collab.], Phys. Rev. **D79**, 052001 (2009) [arXiv:0901.1216].
- P. del Amo Sanchez *et al.* [BABAR Collab.], Phys. Rev. **D82**, 091103 (2010) [arXiv:1008.4080].
- J. Rosner and S. Stone, “Leptonic Decays of Charged Pseudoscalar Leptons,” in this Review.
- P.U.E. Onyisi *et al.* [CLEO Collab.], Phys. Rev. **D79**, 052002 (2009) [arXiv:0901.1147].
- P. Naik *et al.* [CLEO Collab.], Phys. Rev. **D80**, 112004 (2009) [arXiv:0910.3602].
- B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D76**, 052005 (2007) [arXiv:0704.0020].
- LEP  $W$  branching fraction results for this Review of Particle Physics, LEPEWWG/ XSEC/2005-01, <http://lepewwg.web.cern.ch/LEPEWWG/lepww/4f/Winter05/>.

48. P. Abreu *et al.* [DELPHI Collab.], Phys. Lett. **B439**, 209 (1998).
49. I.I.Y. Bigi *et al.*, Phys. Rev. Lett. **71**, 496 (1993) [hep-ph/9304225].
50. A.V. Manohar and M.B. Wise, Phys. Rev. **D49**, 1310 (1994) [hep-ph/9308246].
51. I.I.Y. Bigi *et al.*, Phys. Rev. **D56**, 4017 (1997) [hep-ph/9704245].
52. A.H. Hoang *et al.*, Phys. Rev. **D59**, 074017 (1999) [hep-ph/9811239]; Phys. Rev. Lett. **82**, 277 (1999) [hep-ph/9809423]; A.H. Hoang and T. Teubner, Phys. Rev. **D60**, 114027 (1999) [hep-ph/9904468].
53. N. Isgur and M.B. Wise, Phys. Lett. **B237**, 527 (1990); N. Isgur and M.B. Wise, Phys. Lett. **B232**, 113 (1989).
54. J.P. Lees *et al.* [BABAR Collab.], Phys. Rev. Lett. **109**, 101802 (2012) [arXiv:1205.5442]; M. Huschle *et al.* [Belle Collaboration], Phys. Rev. **D92**, 072014 (2015) [arXiv:1507.03233]; R. Aaij *et al.* [LHCb Collaboration], Phys. Rev. Lett. **115**, 111803 (2015) [arXiv:1506.08614].
55. M. Neubert, Phys. Rev. **D49**, 3392 (1994) [hep-ph/9311325]; Phys. Rev. **D49**, 4623 (1994) [hep-ph/9312311].
56. I.I.Y. Bigi *et al.*, Int. J. Mod. Phys. **A9**, 2467 (1994) [hep-ph/9312359].
57. C.W. Bauer *et al.*, Phys. Lett. **B479**, 395 (2000) [hep-ph/0002161]; Phys. Rev. **D64**, 113004 (2001) [hep-ph/0107074].
58. A. Bornheim *et al.* [CLEO Collab.], Phys. Rev. Lett. **88**, 231803 (2002) [hep-ex/0202019].
59. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D73**, 012006 (2006) [hep-ex/0509040].
60. A. Limosan *et al.* [Belle Collab.], Phys. Lett. **B621**, 28 (2005) [hep-ex/0504046].
61. P. Urquijo *et al.* [Belle Collab.], Phys. Rev. Lett. **104**, 021801 (2010) [arXiv:0907.0379]; J. P. Lees *et al.* [BABAR Collab.], Phys. Rev. **D86**, 032004 (2012) [arXiv:1112.0702].
62. E. Dalgic *et al.*, Phys. Rev. **D73**, 074502 (2006) [Erratum *ibid.* **D75**, 119906 (2007)] [hep-lat/0601021].
63. J.A. Bailey *et al.* [Fermilab Lattice and MILC Collabs.], Phys. Rev. **D92**, 014024 (2015) [arXiv:1503.07839].
64. P. Ball and R. Zwicky, Phys. Rev. **D71**, 014015 (2005) [hep-ph/0406232]; A. Khodjamirian *et al.*, Phys. Rev. **D83**, 094031 (2011) [arXiv:1103.2655].
65. “ $B^\pm$  meson” particle listing, in this *Review*.
66. R. Aaij *et al.* [LHCb Collaboration], Nature Phys. **11**, 743 (2015) [arXiv:1504.01568].
67. H. Albrecht *et al.* [ARGUS Collab.], Phys. Lett. **B192**, 245 (1987).
68. O. Schneider, “ $B^0$ – $\bar{B}^0$  mixing,” in this *Review*.
69. A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **97**, 242003 (2006) [hep-ex/0609040].
70. R. Aaij *et al.* [LHCb Collab.], New J. Phys. **15**, 053021 (2013) [arXiv:1304.4741].
71. Heavy Flavor Averaging Group [22], and updates for Rare Decays at <http://www.slac.stanford.edu/xorg/hfag/rare/index.html>.
72. M. Misiak *et al.*, Phys. Rev. Lett. **114**, 0221801 (2015) [arXiv:1503.01789]; M. Czakon *et al.*, JHEP **1504**, 168 (2015) [arXiv:1503.01791].
73. V. Khachatryan *et al.* [CMS and LHCb Collab.], Nature **522**, 68 (2015) [arXiv:1411.4413].
74. B. Grinstein and D. Pirjol, Phys. Rev. **D62**, 093002 (2000) [hep-ph/0002216]; A. Ali *et al.*, Phys. Lett. **B595**, 323 (2004) [hep-ph/0405075]; M. Beneke *et al.*, Nucl. Phys. **B612**, 25 (2001) [hep-ph/0106067]; S.W. Bosch and G. Buchalla, Nucl. Phys. **B621**, 459 (2002) [hep-ph/0106081]; Z. Ligeti and M. B. Wise, Phys. Rev. **D60**, 117506 (1999) [hep-ph/9905277]; D. Becirevic *et al.*, JHEP **305**, 7 (2003) [hep-lat/0301020]; P. Ball *et al.*, Phys. Rev. **D75**, 054004 (2007) [hep-ph/0612081]; W. Wang *et al.*, arXiv:0711.0432; C.D. Lu *et al.*, Phys. Rev. **D76**, 014013 (2007) [hep-ph/0701265].
75. A. J. Buras *et al.*, Phys. Rev. Lett. **95**, 261805 (2005) [hep-ph/0508165].
76. A.V. Artamonov *et al.* [E949 Collab.], Phys. Rev. Lett. **101**, 191802 (2008) [arXiv:0808.2459]; Phys. Rev. **D79**, 092004 (2009) [arXiv:0903.0030].
77. D. Acosta *et al.* [CDF Collab.], Phys. Rev. Lett. **95**, 102002 (2005) [hep-ex/0505091].
78. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. Lett. **107**, 121802 (2011) [arXiv:1106.5436].
79. CMS-PAS-TOP-11-029 (2011) [CMS Collab.].
80. T. A. Aaltonen *et al.* [CDF and DØ Collab.], Phys. Rev. Lett. **115**, 152003 (2015) [arXiv:1503.05027].
81. LHC Top Working group, [https://lpsc.web.cern.ch/lpsc/index.php?page=top\\_wg](https://lpsc.web.cern.ch/lpsc/index.php?page=top_wg), WG Plots: Single Top Quark Production, November 2015; in addition, a new t-channel result at 13 TeV, ATLAS-CONF-2015-079 is included.
82. J. Swain and L. Taylor, Phys. Rev. **D58**, 093006 (1998) [hep-ph/9712420].
83. “ $K_L^0$  meson” particle listing, in this *Review*.
84. A. J. Buras, D. Guadagnoli and G. Isidori, Phys. Lett. **B688**, 309 (2010) [arXiv:1002.3612]; for earlier discussions, see [85].
85. E. A. Andriyash, G. G. Ovanesyan and M. I. Vysotsky, Phys. Lett. **B599**, 253 (2004) [hep-ph/0310314]; K. Anikeev *et al.*, hep-ph/0201071; A. J. Buras and D. Guadagnoli, Phys. Rev. **D78**, 033005 (2008) [arXiv:0805.3887].
86. T. Inami and C.S. Lim, Prog. Theor. Phys. **65**, 297 (1981) [Erratum *ibid.* **65**, 1772 (1981)].
87. J.M. Flynn and L. Randall, Phys. Lett. **B224**, 221 (1989); G. Buchalla, A. J. Buras, and M. K. Harlander, Nucl. Phys. **B337**, 313 (1990).
88. M. Ciuchini *et al.*, Phys. Lett. **B301**, 263 (1993) [hep-ph/9212203]; A.J. Buras, M. Jamin, and M.E. Lautenbacher, Nucl. Phys. **B408**, 209 (1993) [hep-ph/9303284].
89. T. Hambye *et al.*, Nucl. Phys. **B564**, 391 (2000) [hep-ph/9906434].
90. S. Bertolini *et al.*, Phys. Rev. **D63**, 056009 (2001) [hep-ph/0002234].
91. A. Pich, hep-ph/0410215.
92. Z. Bai *et al.* [RBC and UKQCD Collabs], Phys. Rev. Lett. **115**, 212001 (2015) [arXiv:1505.07863].
93. A.B. Carter and A.I. Sanda, Phys. Rev. Lett. **45**, 952 (1980); Phys. Rev. **D23**, 1567 (1981).
94. A more detailed discussion and references can be found in: D. Kirkby and Y. Nir, “ $CP$  violation in meson decays,” in this *Review*.
95. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D79**, 072009 (2009) [arXiv:0902.1708].
96. I. Adachi *et al.* [Belle Collab.], Phys. Rev. Lett. **108**, 171802 (2012) [arXiv:1201.4643].
97. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **115**, 031601 (2015) [arXiv:1503.07089].
98. Heavy Flavor Averaging Group [22], Summer 2015 updates for Unitarity Triangle Parameters: <http://www.slac.stanford.edu/xorg/hfag/triangle/summer2015/>.
99. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D71**, 032005 (2005) [hep-ex/0411016].
100. R. Itoh *et al.* [Belle Collab.], Phys. Rev. Lett. **95**, 091601 (2005) [hep-ex/0504030].
101. P. Krokovny *et al.* [Belle Collab.], Phys. Rev. Lett. **97**, 081801 (2006) [hep-ex/0507065].
102. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. Lett. **99**, 231802 (2007) [arXiv:0708.1544].
103. A. Abdesselam *et al.* [BABAR and Belle Collab.], Phys. Rev. Lett. **115**, 121604 (2015) [arXiv:1505.04147].



104. M. Gronau and D. London, Phys. Rev. Lett. **65**, 3381 (1990).
105. A. Höcker *et al.*, Eur. Phys. J. **C21**, 225 (2001) [hep-ph/0104062]; see also Ref. [6] and updates at <http://ckmfitter.in2p3.fr/>. The “EPS-HEP 2015” results are used in this article.
106. J. Zhang *et al.* [Belle Collab.], Phys. Rev. Lett. **91**, 221801 (2003) [hep-ex/0306007];  
A. Somov *et al.* [Belle Collab.], Phys. Rev. Lett. **96**, 171801 (2006) [hep-ex/0601024];  
B. Aubert *et al.* [BABAR Collab.], Phys. Rev. Lett. **97**, 261801 (2006) [hep-ex/0607092]; Phys. Rev. **D76**, 052007 (2007) [arXiv:0705.2157].
107. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D78**, 071104 (2008) [arXiv:0807.4977].
108. A.F. Falk *et al.*, Phys. Rev. **D69**, 011502 (2004) [hep-ph/0310242].
109. H.R. Quinn and A.E. Snyder, Phys. Rev. **D48**, 2139 (1993).
110. A. Kusaka *et al.* [Belle Collab.], Phys. Rev. Lett. **98**, 221602 (2007) [hep-ex/0701015].
111. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D88**, 121003 (2013) [arXiv:1304.3503].
112. M. Bona *et al.* [UTfit Collab.], JHEP **507**, 28 (2005) [hep-ph/0501199], and updates at <http://www.utfit.org/>.
113. M. Gronau and D. London, Phys. Lett. **B253**, 483 (1991).
114. M. Gronau and D. Wyler, Phys. Lett. **B265**, 172 (1991).
115. D. Atwood *et al.*, Phys. Rev. Lett. **78**, 3257 (1997) [hep-ph/9612433]; Phys. Rev. **D63**, 036005 (2001) [hep-ph/0008090].
116. A. Bondar, talk at the Belle analysis workshop, Novosibirsk, September 2002;  
A. Poluektov *et al.* [Belle Collab.], Phys. Rev. **D70**, 072003 (2004) [hep-ex/0406067].
117. A. Giri *et al.*, Phys. Rev. **D68**, 054018 (2003) [hep-ph/0303187].
118. A. Poluektov *et al.* [Belle Collab.], Phys. Rev. **D81**, 112002 (2010) [arXiv:1003.3360].
119. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. Lett. **105**, 121801 (2010) [arXiv:1005.1096].
120. R. Aaij *et al.*, JHEP **1410**, 097 (2014) [arXiv:1408.2748].
121. Y. Grossman *et al.*, Phys. Rev. **D72**, 031501 (2005) [hep-ph/0505270].
122. A. Amorim *et al.*, Phys. Rev. **D59**, 056001 (1999) [hep-ph/9807364].
123. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D71**, 112003 (2005) [hep-ex/0504035]; Phys. Rev. **D73**, 111101 (2006) [hep-ex/0602049]; F.J. Ronga *et al.* [Belle Collab.], Phys. Rev. **D73**, 092003 (2006) [hep-ex/0604013]; S. Bahinipati *et al.* [Belle Collab.], Phys. Rev. **D84**, 021101 (2011) [arXiv:1102.0888].
124. R. Aleksan *et al.*, Z. Phys. **C54**, 653 (1992).
125. R. Aaij *et al.*, JHEP **1411**, 060 (2014) [arXiv:1407.6127].
126. G.P. Dubois-Felsmann *et al.*, hep-ph/0308262; G. Eigen *et al.*, Phys. Rev. **D89**, 033004 (2014) [arXiv:1301.5867].
127. “The BABAR physics book: Physics at an asymmetric B factory,” (P.F. Harrison and H.R. Quinn, eds.), SLAC-R-0504, 1998.
128. S. Plaszczynski and M.H. Schune, hep-ph/9911280.
129. M. Bona *et al.* [UTfit Collab.], JHEP **0803**, 049 (2008) [arXiv:0707.0636].
130. We thank the CKMfitter and UTfit groups for performing fits and preparing plots using input values from this *Review*.
131. H.E. Haber, Nucl. Phys. Proc. Supp. **62**, 469 (1998) [hep-ph/9709450];  
Y. Nir, hep-ph/0109090.
132. J.M. Soares and L. Wolfenstein, Phys. Rev. **D47**, 1021 (1993);  
T. Goto *et al.*, Phys. Rev. **D53**, 6662 (1996) [hep-ph/9506311];  
J.P. Silva and L. Wolfenstein, Phys. Rev. **D55**, 5331 (1997) [hep-ph/9610208].
133. Y. Grossman, Z. Ligeti, and Y. Nir, Prog. Theor. Phys. **122**, 125 (2009) [arXiv:0904.4262].
134. G. Isidori, Y. Nir, and G. Perez, Ann. Rev. Nucl. and Part. Sci. **60**, 355 (2010) [arXiv:1002.0900].
135. Z. Ligeti, Int. J. Mod. Phys. **A20**, 5105 (2005) [hep-ph/0408267].
136. J. Charles *et al.*, arXiv:1309.2293.
137. K. Agashe *et al.*, hep-ph/0509117.
138. S. Laplace *et al.*, Phys. Rev. **D65**, 094040 (2002) [hep-ph/0202010].
139. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D89**, 012002 (2014) [arXiv:1310.0447].
140. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **114**, 1041801 (2015) [arXiv:1411.3104].
141. G. Buchalla and A.J. Buras, Phys. Lett. **B333**, 221 (1994) [hep-ph/9405259].



### 13. *CP* VIOLATION IN THE QUARK SECTOR

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The *CP* transformation combines charge conjugation *C* with parity *P*. Under *C*, particles and antiparticles are interchanged, by conjugating all internal quantum numbers, *e.g.*,  $Q \rightarrow -Q$  for electromagnetic charge. Under *P*, the handedness of space is reversed,  $\vec{x} \rightarrow -\vec{x}$ . Thus, for example, a left-handed electron  $e_L^-$  is transformed under *CP* into a right-handed positron,  $e_R^+$ .

If *CP* were an exact symmetry, the laws of Nature would be the same for matter and for antimatter. We observe that most phenomena are *C*- and *P*-symmetric, and therefore, also *CP*-symmetric. In particular, these symmetries are respected by the gravitational, electromagnetic, and strong interactions. The weak interactions, on the other hand, violate *C* and *P* in the strongest possible way. For example, the charged *W* bosons couple to left-handed electrons,  $e_L^-$ , and to their *CP*-conjugate right-handed positrons,  $e_R^+$ , but to neither their *C*-conjugate left-handed positrons,  $e_L^+$ , nor their *P*-conjugate right-handed electrons,  $e_R^-$ . While weak interactions violate *C* and *P* separately, *CP* is still preserved in most weak interaction processes. The *CP* symmetry is, however, violated in certain rare processes, as discovered in neutral *K* decays in 1964 [1], and observed in recent years in *B* decays. A  $K_L$  meson decays more often to  $\pi^- e^+ \nu_e$  than to  $\pi^+ e^- \bar{\nu}_e$ , thus allowing electrons and positrons to be unambiguously distinguished, but the decay-rate asymmetry is only at the 0.003 level. The *CP*-violating effects observed in the *B* system are larger: the parameter describing the *CP* asymmetry in the decay time distribution of  $B^0/\bar{B}^0$  meson transitions to *CP* eigenstates like  $J/\psi K_S$  is about 0.7 [2,3]. These effects are related to  $K^0-\bar{K}^0$  and  $B^0-\bar{B}^0$  mixing, but *CP* violation arising solely from decay amplitudes has also been observed, first in  $K \rightarrow \pi\pi$  decays [4–6], and more recently in  $B^0$  [7,8],  $B^+$  [9–11], and  $B_s^0$  [12] decays. Similar effects could also occur, but have not yet been observed, in decays of *b* baryons. *CP* violation is not yet experimentally established in the *D* system, where the Standard Model effects are expected to be  $\mathcal{O}(10^{-3})$ . Moreover, *CP* violation has not yet been observed in processes involving the top quark, nor in flavor-conserving processes such as electric dipole moments, nor in the lepton sector; for all of these any significant observation would be a clear indication of physics beyond the Standard Model.

In addition to parity and to continuous Lorentz transformations, there is one other spacetime operation that could be a symmetry of the interactions: time reversal *T*,  $t \rightarrow -t$ . Violations of *T* symmetry have been observed in neutral *K* decays [13]. More recently, exploiting the fact that for neutral *B* mesons both flavor tagging and *CP* tagging can be used [14], *T* violation has been observed between states that are not *CP*-conjugate [15]. Moreover, *T* violation is expected as a corollary of *CP* violation if the combined *CPT* transformation is a fundamental symmetry of Nature [16]. All observations indicate that *CPT* is indeed a symmetry of Nature. Furthermore, one cannot build a locally Lorentz-invariant quantum field theory with a Hermitian Hamiltonian that violates *CPT*. (At several points in our discussion, we avoid assumptions about *CPT*, in order to identify cases where evidence for *CP* violation relies on assumptions about *CPT*.)

Within the Standard Model, *CP* symmetry is broken by complex phases in the Yukawa couplings (that is, the couplings of the Higgs scalar to quarks). When all manipulations to remove unphysical phases in this model are exhausted, one finds that there is a single *CP*-violating parameter [17]. In the basis of mass eigenstates, this single phase appears in the  $3 \times 3$  unitary matrix that gives the *W*-boson couplings to an up-type antiquark and a down-type quark. (If the Standard Model is supplemented with Majorana mass terms for the neutrinos, the analogous mixing matrix for leptons has three *CP*-violating phases.) The beautifully consistent and economical Standard-Model description of *CP* violation in terms of Yukawa couplings, known as the Kobayashi-Maskawa (KM) mechanism [17], agrees with all measurements to date. (Some measurements are in tension with the predictions, and are discussed in more detail below. Pending verification, the results are not considered to change the overall picture of agreement with the Standard Model.) Furthermore, one can fit the data allowing new physics contributions to loop processes to compete with, or even dominate over, the Standard Model

amplitudes [18,19]. Such an analysis provides model-independent proof that the KM phase is different from zero, and that the matrix of three-generation quark mixing is the dominant source of *CP* violation in meson decays.

The current level of experimental accuracy and the theoretical uncertainties involved in the interpretation of the various observations leave room, however, for additional subdominant sources of *CP* violation from new physics. Indeed, almost all extensions of the Standard Model imply that there are such additional sources. Moreover, *CP* violation is a necessary condition for baryogenesis, the process of dynamically generating the matter-antimatter asymmetry of the Universe [20]. Despite the phenomenological success of the KM mechanism, it fails (by several orders of magnitude) to accommodate the observed asymmetry [21]. This discrepancy strongly suggests that Nature provides additional sources of *CP* violation beyond the KM mechanism. The evidence for neutrino masses implies that *CP* can be violated also in the lepton sector. This situation makes leptogenesis [22,23], a scenario where *CP*-violating phases in the Yukawa couplings of the neutrinos play a crucial role in the generation of the baryon asymmetry, a very attractive possibility. The expectation of new sources motivates the large ongoing experimental effort to find deviations from the predictions of the KM mechanism.

*CP* violation can be experimentally searched for in a variety of processes, such as hadron decays, electric dipole moments of neutrons, electrons and nuclei, and neutrino oscillations. Hadron decays via the weak interaction probe flavor-changing *CP* violation. The search for electric dipole moments may find (or constrain) sources of *CP* violation that, unlike the KM phase, are not related to flavor-changing couplings. Following the discovery of the Higgs boson [24,25], searches for *CP* violation in the Higgs sector are becoming feasible. Future searches for *CP* violation in neutrino oscillations might provide further input on leptogenesis.

The present measurements of *CP* asymmetries provide some of the strongest constraints on the weak couplings of quarks. Future measurements of *CP* violation in *K*, *D*, *B*, and  $B_s^0$  meson decays will provide additional constraints on the flavor parameters of the Standard Model, and can probe new physics. In this review, we give the formalism and basic physics that are relevant to present and near future measurements of *CP* violation in the quark sector.

Before going into details, we list here the observables where *CP* violation has been observed at a level above  $5\sigma$  [26–28]:

- Indirect *CP* violation in  $K \rightarrow \pi\pi$  and  $K \rightarrow \pi\ell\nu$  decays, and in the  $K_L \rightarrow \pi^+\pi^-e^+e^-$  decay, is given by

$$|\epsilon| = (2.228 \pm 0.011) \times 10^{-3}. \quad (13.1)$$

- Direct *CP* violation in  $K \rightarrow \pi\pi$  decays is given by

$$\mathcal{R}e(\epsilon'/\epsilon) = (1.65 \pm 0.26) \times 10^{-3}. \quad (13.2)$$

- *CP* violation in the interference of mixing and decay in the tree-dominated  $b \rightarrow c\bar{c}s$  transitions, such as  $B^0 \rightarrow \psi K^0$ , is given by (we use  $K^0$  throughout to denote results that combine  $K_S$  and  $K_L$  modes, but use the sign appropriate to  $K_S$ ):

$$S_{\psi K^0} = +0.691 \pm 0.017. \quad (13.3)$$

- *CP* violation in the interference of mixing and decay in modes governed by the tree-dominated  $b \rightarrow c\bar{u}d$  transitions is given by

$$S_{D_{CP}^{(*)}h^0} = +0.63 \pm 0.11, \quad (13.4)$$

- *CP* violation in the interference of mixing and decay in various modes related to  $b \rightarrow c\bar{u}d$  transitions is given by

$$S_{\psi\pi^0} = -0.93 \pm 0.15, \quad (13.5)$$

$$S_{D^+D^-} = -0.98 \pm 0.17. \quad (13.6)$$

$$S_{D^{*+}D^{*-}} = -0.71 \pm 0.09. \quad (13.7)$$

- *CP* violation in the interference of mixing and decay in various modes related to  $b \rightarrow q\bar{q}s$  (penguin) transitions is given by

$$S_{\phi K^0} = +0.74_{-0.13}^{+0.11}, \quad (13.8)$$

$$S_{\eta' K^0} = +0.63 \pm 0.06, \quad (13.9)$$

$$S_{f_0 K^0} = +0.69_{-0.12}^{+0.10}, \quad (13.10)$$

$$S_{K+K-K_S} = +0.68_{-0.10}^{+0.09}, \quad (13.11)$$

- *CP* violation in the interference of mixing and decay in the  $B^0 \rightarrow \pi^+\pi^-$  mode is given by

$$S_{\pi^+\pi^-} = -0.66 \pm 0.06. \quad (13.12)$$

- Direct *CP* violation in the  $B^0 \rightarrow \pi^+\pi^-$  mode is given by

$$C_{\pi^+\pi^-} = -0.31 \pm 0.05. \quad (13.13)$$

- Direct *CP* violation in the  $\bar{B}^0 \rightarrow K^-\pi^+$  mode is given by

$$\mathcal{A}_{\bar{B}^0 \rightarrow K^-\pi^+} = -0.082 \pm 0.006. \quad (13.14)$$

- Direct *CP* violation in  $B^+ \rightarrow D_+K^+$  decays ( $D_+$  is the *CP*-even neutral  $D$  state) is given by

$$\mathcal{A}_{B^+ \rightarrow D_+K^+} = +0.195 \pm 0.027. \quad (13.15)$$

- Direct *CP* violation in the  $\bar{B}_s^0 \rightarrow K^+\pi^-$  mode is given by

$$\mathcal{A}_{\bar{B}_s^0 \rightarrow K^+\pi^-} = +0.26 \pm 0.04. \quad (13.16)$$

- Direct *CP* violation in  $B^+ \rightarrow K^+K^-\pi^+$  decays is given by

$$\mathcal{A}_{B^+ \rightarrow K^+K^-\pi^+} = -0.118 \pm 0.022. \quad (13.17)$$

In addition, large *CP* violation effects have recently been observed in certain regions of the phase space of  $B^+ \rightarrow K^+K^-K^+$ ,  $\pi^+\pi^-K^+$ ,  $\pi^+\pi^-\pi^+$  and  $K^+K^-\pi^+$  decays.

### 13.1. Formalism

The phenomenology of *CP* violation for neutral flavored mesons is particularly interesting, since many of the observables can be cleanly interpreted. Although the phenomenology is superficially different for  $K^0$ ,  $D^0$ ,  $B^0$ , and  $B_s^0$  decays, this is primarily because each of these systems is governed by a different balance between decay rates, oscillations, and lifetime splitting. However, the general considerations presented in this section are identical for all flavored neutral pseudoscalar mesons. The phenomenology of *CP* violation for neutral mesons that do not carry flavor quantum numbers (such as the  $\eta^{(\prime)}$  state) is quite different: such states are their own antiparticles and have definite *CP* eigenvalues, so the signature of *CP* violation is simply the decay to a final state with the opposite *CP*. Such decays are mediated by the electromagnetic or (OZI-suppressed) strong interaction, where *CP* violation is not expected and has not yet been observed. In the remainder of this review, we restrict ourselves to considerations of weakly decaying hadrons.

In this section, we present a general formalism for, and classification of, *CP* violation in the decay of a weakly decaying hadron, denoted  $M$ . We pay particular attention to the case that  $M$  is a  $K^0$ ,  $D^0$ ,  $B^0$ , or  $B_s^0$  meson. Subsequent sections describe the *CP*-violating phenomenology, approximations, and alternative formalisms that are specific to each system.

#### 13.1.1. Charged- and neutral-hadron decays :

We define decay amplitudes of  $M$  (which could be charged or neutral) and its *CP* conjugate  $\bar{M}$  to a multi-particle final state  $f$  and its *CP* conjugate  $\bar{f}$  as

$$A_f = \langle f | \mathcal{H} | M \rangle, \quad \bar{A}_{\bar{f}} = \langle \bar{f} | \mathcal{H} | \bar{M} \rangle, \\ \bar{A}_{\bar{f}} = \langle \bar{f} | \mathcal{H} | M \rangle, \quad \bar{A}_{\bar{f}} = \langle \bar{f} | \mathcal{H} | \bar{M} \rangle, \quad (13.18)$$

where  $\mathcal{H}$  is the Hamiltonian governing weak interactions. The action of *CP* on these states introduces phases  $\xi_M$  and  $\xi_f$  that depend on their flavor content, according to

$$CP|M\rangle = e^{+i\xi_M} |\bar{M}\rangle, \quad CP|f\rangle = e^{+i\xi_f} |\bar{f}\rangle, \quad (13.19)$$

with

$$CP|\bar{M}\rangle = e^{-i\xi_M} |M\rangle, \quad CP|\bar{f}\rangle = e^{-i\xi_f} |f\rangle \quad (13.20)$$

so that  $(CP)^2 = 1$ . The phases  $\xi_M$  and  $\xi_f$  are arbitrary and unobservable because of the flavor symmetry of the strong interaction. If *CP* is conserved by the dynamics,  $[CP, \mathcal{H}] = 0$ , then  $A_f$  and  $\bar{A}_{\bar{f}}$  have the same magnitude and an arbitrary unphysical relative phase

$$\bar{A}_{\bar{f}} = e^{i(\xi_f - \xi_M)} A_f. \quad (13.21)$$

#### 13.1.2. Neutral-meson mixing :

A state that is initially a superposition of  $M^0$  and  $\bar{M}^0$ , say

$$|\psi(0)\rangle = a(0)|M^0\rangle + b(0)|\bar{M}^0\rangle, \quad (13.22)$$

will evolve in time acquiring components that describe all possible decay final states  $\{f_1, f_2, \dots\}$ , that is,

$$|\psi(t)\rangle = a(t)|M^0\rangle + b(t)|\bar{M}^0\rangle + c_1(t)|f_1\rangle + c_2(t)|f_2\rangle + \dots \quad (13.23)$$

If we are interested in computing only the values of  $a(t)$  and  $b(t)$  (and not the values of all  $c_i(t)$ ), and if the times  $t$  in which we are interested are much larger than the typical strong interaction scale, then we can use a much simplified formalism [29]. The simplified time evolution is determined by a  $2 \times 2$  effective Hamiltonian  $\mathbf{H}$  that is not Hermitian, since otherwise the mesons would only oscillate and not decay. Any complex matrix, such as  $\mathbf{H}$ , can be written in terms of Hermitian matrices  $\mathbf{M}$  and  $\mathbf{\Gamma}$  as

$$\mathbf{H} = \mathbf{M} - \frac{i}{2} \mathbf{\Gamma}. \quad (13.24)$$

$\mathbf{M}$  and  $\mathbf{\Gamma}$  are associated with  $(M^0, \bar{M}^0) \leftrightarrow (M^0, \bar{M}^0)$  transitions via off-shell (dispersive), and on-shell (absorptive) intermediate states, respectively. Diagonal elements of  $\mathbf{M}$  and  $\mathbf{\Gamma}$  are associated with the flavor-conserving transitions  $M^0 \rightarrow M^0$  and  $\bar{M}^0 \rightarrow \bar{M}^0$ , while off-diagonal elements are associated with flavor-changing transitions  $M^0 \leftrightarrow \bar{M}^0$ .

The eigenvectors of  $\mathbf{H}$  have well-defined masses and decay widths. To specify the components of the strong interaction eigenstates,  $M^0$  and  $\bar{M}^0$ , in the light ( $M_L$ ) and heavy ( $M_H$ ) mass eigenstates, we introduce three complex parameters:  $p$ ,  $q$ , and, for the case that both *CP* and *CPT* are violated in mixing,  $z$ :

$$|M_L\rangle \propto p\sqrt{1-z}|M^0\rangle + q\sqrt{1+z}|\bar{M}^0\rangle \\ |M_H\rangle \propto p\sqrt{1+z}|M^0\rangle - q\sqrt{1-z}|\bar{M}^0\rangle, \quad (13.25)$$

with the normalization  $|q|^2 + |p|^2 = 1$  when  $z = 0$ . (Another possible choice, which is in standard usage for  $K$  mesons, defines the mass eigenstates according to their lifetimes:  $K_S$  for the short-lived and  $K_L$  for the long-lived state. The  $K_L$  is experimentally found to be the heavier state. Yet another choice is often used for the  $D$  mesons [30]: the eigenstates are labelled according to their dominant *CP* content.)

The real and imaginary parts of the eigenvalues  $\omega_{L,H}$  corresponding to  $|M_{L,H}\rangle$  represent their masses and decay widths, respectively. The mass and width splittings are

$$\Delta m \equiv m_H - m_L = \mathcal{R}e(\omega_H - \omega_L), \\ \Delta \Gamma \equiv \Gamma_H - \Gamma_L = -2\mathcal{I}m(\omega_H - \omega_L). \quad (13.26)$$

Note that here  $\Delta m$  is positive by definition, while the sign of  $\Delta\Gamma$  must be experimentally determined. The sign of  $\Delta\Gamma$  has not yet been established for  $B^0$  mesons, while  $\Delta\Gamma < 0$  is established for  $K$  and  $B_s^0$  mesons. The Standard Model predicts  $\Delta\Gamma < 0$  also for  $B_{(s)}^0$  mesons (for this reason,  $\Delta\Gamma = \Gamma_L - \Gamma_H$ , which is still a signed quantity, is often used in the  $B^0$  and  $B_s^0$  literature and is the convention used in the PDG experimental summaries).

Solving the eigenvalue problem for  $\mathbf{H}$  yields

$$\left(\frac{q}{p}\right)^2 = \frac{\mathbf{M}_{12}^* - (i/2)\mathbf{\Gamma}_{12}^*}{\mathbf{M}_{12} - (i/2)\mathbf{\Gamma}_{12}} \quad (13.27)$$

and

$$z \equiv \frac{\delta m - (i/2)\delta\Gamma}{\Delta m - (i/2)\Delta\Gamma}, \quad (13.28)$$

where

$$\delta m \equiv \mathbf{M}_{11} - \mathbf{M}_{22}, \quad \delta\Gamma \equiv \mathbf{\Gamma}_{11} - \mathbf{\Gamma}_{22} \quad (13.29)$$

are the differences in effective mass and decay-rate expectation values for the strong interaction states  $M^0$  and  $\bar{M}^0$ .

If either  $CP$  or  $CPT$  is a symmetry of  $\mathbf{H}$  (independently of whether  $T$  is conserved or violated), then the values of  $\delta m$  and  $\delta\Gamma$  are both zero, and hence  $z = 0$ . We also find that

$$\omega_H - \omega_L = 2\sqrt{\left(\mathbf{M}_{12} - \frac{i}{2}\mathbf{\Gamma}_{12}\right)\left(\mathbf{M}_{12}^* - \frac{i}{2}\mathbf{\Gamma}_{12}^*\right)}. \quad (13.30)$$

If either  $CP$  or  $T$  is a symmetry of  $\mathbf{H}$  (independently of whether  $CPT$  is conserved or violated), then  $\mathbf{\Gamma}_{12}/\mathbf{M}_{12}$  is real, leading to

$$\left(\frac{q}{p}\right)^2 = e^{2i\xi_M} \Rightarrow \left|\frac{q}{p}\right| = 1, \quad (13.31)$$

where  $\xi_M$  is the arbitrary unphysical phase introduced in Eq. (13.20). If, and only if,  $CP$  is a symmetry of  $\mathbf{H}$  (independently of  $CPT$  and  $T$ ), then both of the above conditions hold, with the result that the mass eigenstates are orthogonal

$$\langle M_H | M_L \rangle = |p|^2 - |q|^2 = 0. \quad (13.32)$$

### 13.1.3. CP-violating observables :

All  $CP$ -violating observables in  $M$  and  $\bar{M}$  decays to final states  $f$  and  $\bar{f}$  can be expressed in terms of phase-convention-independent combinations of  $A_f$ ,  $\bar{A}_f$ ,  $A_{\bar{f}}$ , and  $\bar{A}_{\bar{f}}$ , together with, for neutral meson decays only,  $q/p$ .  $CP$  violation in charged meson and all baryon decays depends only on the combination  $|\bar{A}_{\bar{f}}/A_f|$ , while  $CP$  violation in flavored neutral meson decays is complicated by  $M^0 \leftrightarrow \bar{M}^0$  oscillations, and depends, additionally, on  $|q/p|$  and on  $\lambda_f \equiv (q/p)(\bar{A}_f/A_f)$ .

The decay rates of the two neutral kaon mass eigenstates,  $K_S$  and  $K_L$ , are different enough ( $\Gamma_S/\Gamma_L \sim 500$ ) that one can, in most cases, actually study their decays independently. For  $D^0$ ,  $B^0$ , and  $B_s^0$  mesons, however, values of  $\Delta\Gamma/\Gamma$  (where  $\Gamma \equiv (\Gamma_H + \Gamma_L)/2$ ) are relatively small, and so both mass eigenstates must be considered in their evolution. We denote the state of an initially pure  $|M^0\rangle$  or  $|\bar{M}^0\rangle$  after an elapsed proper time  $t$  as  $|M_{\text{phys}}^0(t)\rangle$  or  $|\bar{M}_{\text{phys}}^0(t)\rangle$ , respectively. Using the effective Hamiltonian approximation, but not assuming  $CPT$  is a good symmetry, we obtain

$$\begin{aligned} |M_{\text{phys}}^0(t)\rangle &= (g_+(t) + z g_-(t)) |M^0\rangle - \sqrt{1-z^2} \frac{q}{p} g_-(t) |\bar{M}^0\rangle, \\ |\bar{M}_{\text{phys}}^0(t)\rangle &= (g_+(t) - z g_-(t)) |\bar{M}^0\rangle - \sqrt{1-z^2} \frac{p}{q} g_-(t) |M^0\rangle, \end{aligned} \quad (13.33)$$

where

$$g_{\pm}(t) \equiv \frac{1}{2} \left( e^{-im_H t - \frac{1}{2}\Gamma_H t} \pm e^{-im_L t - \frac{1}{2}\Gamma_L t} \right) \quad (13.34)$$

and  $z = 0$  if either  $CPT$  or  $CP$  is conserved.

Defining  $x \equiv \Delta m/\Gamma$  and  $y \equiv \Delta\Gamma/(2\Gamma)$ , and assuming  $z = 0$ , one obtains the following time-dependent decay rates:

$$\begin{aligned} \frac{d\Gamma[M_{\text{phys}}^0(t) \rightarrow f]/dt}{e^{-\Gamma t} \mathcal{N}_f} &= \\ & \left( |A_f|^2 + |(q/p)\bar{A}_f|^2 \right) \cosh(y\Gamma t) + \left( |A_f|^2 - |(q/p)\bar{A}_f|^2 \right) \cos(x\Gamma t) \\ & + 2\mathcal{R}e((q/p)A_f^* \bar{A}_f) \sinh(y\Gamma t) - 2\mathcal{I}m((q/p)A_f^* \bar{A}_f) \sin(x\Gamma t), \end{aligned} \quad (13.35)$$

$$\begin{aligned} \frac{d\Gamma[\bar{M}_{\text{phys}}^0(t) \rightarrow \bar{f}]/dt}{e^{-\Gamma t} \mathcal{N}_{\bar{f}}} &= \\ & \left( |(p/q)A_f|^2 + |\bar{A}_f|^2 \right) \cosh(y\Gamma t) - \left( |(p/q)A_f|^2 - |\bar{A}_f|^2 \right) \cos(x\Gamma t) \\ & + 2\mathcal{R}e((p/q)A_f \bar{A}_f^*) \sinh(y\Gamma t) - 2\mathcal{I}m((p/q)A_f \bar{A}_f^*) \sin(x\Gamma t), \end{aligned} \quad (13.36)$$

where  $\mathcal{N}_f$  is a common, time-independent, normalization factor that can be determined bearing in mind that the range of  $t$  is  $0 < t < \infty$ . Decay rates to the  $CP$ -conjugate final state  $\bar{f}$  are obtained analogously, with  $\mathcal{N}_f = \mathcal{N}_{\bar{f}}$  and the substitutions  $A_f \rightarrow A_{\bar{f}}$  and  $\bar{A}_f \rightarrow \bar{A}_{\bar{f}}$  in Eqs. (13.35, 13.36). Terms proportional to  $|A_f|^2$  or  $|\bar{A}_f|^2$  are associated with decays that occur without any net  $M^0 \leftrightarrow \bar{M}^0$  oscillation, while terms proportional to  $|(q/p)\bar{A}_f|^2$  or  $|(p/q)A_f|^2$  are associated with decays following a net oscillation. The  $\sinh(y\Gamma t)$  and  $\sin(x\Gamma t)$  terms of Eqs. (13.35, 13.36) are associated with the interference between these two cases. Note that, in multi-body decays, amplitudes are functions of phase-space variables. Interference may be present in some regions but not others, and is strongly influenced by resonant substructure.

When neutral pseudoscalar mesons are produced coherently in pairs from the decay of a vector resonance,  $V \rightarrow M^0 \bar{M}^0$  (for example,  $\Upsilon(4S) \rightarrow B^0 \bar{B}^0$  or  $\phi \rightarrow K^0 \bar{K}^0$ ), the time-dependence of their subsequent decays to final states  $f_1$  and  $f_2$  has a similar form to Eqs. (13.35, 13.36):

$$\begin{aligned} \frac{d\Gamma[V_{\text{phys}}(t_1, t_2) \rightarrow f_1 f_2]/d(\Delta t)}{e^{-\Gamma|\Delta t|} \mathcal{N}_{f_1 f_2}} &= \\ & \left( |a_+|^2 + |a_-|^2 \right) \cosh(y\Gamma \Delta t) + \left( |a_+|^2 - |a_-|^2 \right) \cos(x\Gamma \Delta t) \\ & - 2\mathcal{R}e(a_+^* a_-) \sinh(y\Gamma \Delta t) + 2\mathcal{I}m(a_+^* a_-) \sin(x\Gamma \Delta t), \end{aligned} \quad (13.37)$$

where  $\Delta t \equiv t_2 - t_1$  is the difference in the production times,  $t_1$  and  $t_2$ , of  $f_1$  and  $f_2$ , respectively, and the dependence on the average decay time and on decay angles has been integrated out. The normalisation factor  $\mathcal{N}_{f_1 f_2}$  can be evaluated, noting that the range of  $\Delta t$  is  $-\infty < \Delta t < \infty$ . The coefficients in Eq. (13.37) are determined by the amplitudes for no net oscillation from  $t_1 \rightarrow t_2$ ,  $\bar{A}_{f_1} \bar{A}_{f_2}$ , and  $A_{f_1} \bar{A}_{f_2}$ , and for a net oscillation,  $(q/p)\bar{A}_{f_1} \bar{A}_{f_2}$  and  $(p/q)A_{f_1} \bar{A}_{f_2}$ , via

$$\begin{aligned} a_+ &\equiv \bar{A}_{f_1} A_{f_2} - A_{f_1} \bar{A}_{f_2}, \\ a_- &\equiv -\sqrt{1-z^2} \left( \frac{q}{p} \bar{A}_{f_1} \bar{A}_{f_2} - \frac{p}{q} A_{f_1} A_{f_2} \right) + z (\bar{A}_{f_1} A_{f_2} + A_{f_1} \bar{A}_{f_2}). \end{aligned} \quad (13.38)$$

Assuming  $CPT$  conservation,  $z = 0$ , and identifying  $\Delta t \rightarrow t$  and  $f_2 \rightarrow f$ , we find that Eqs. (13.37, 13.38) reduce essentially to Eq. (13.35) with  $A_{f_1} = 0$ ,  $\bar{A}_{f_1} = 1$ , or to Eq. (13.36) with  $\bar{A}_{f_1} = 0$ ,  $A_{f_1} = 1$ . Indeed, such a situation plays an important role in experiments that exploit the coherence of  $V \rightarrow M^0 \bar{M}^0$  (for example  $\psi(3770) \rightarrow D^0 \bar{D}^0$  or  $\Upsilon(4S) \rightarrow B^0 \bar{B}^0$ ) production. Final states  $f_1$  with  $A_{f_1} = 0$  or  $\bar{A}_{f_1} = 0$  are called tagging states, because they identify the decaying pseudoscalar meson as, respectively,  $\bar{M}^0$  or  $M^0$ . Before one of  $M^0$  or  $\bar{M}^0$  decays, they evolve in phase, so that there is always one  $M^0$  and one  $\bar{M}^0$  present. A tagging decay of one meson sets the clock for the time evolution of the other: it starts at  $t_1$  as purely  $M^0$  or  $\bar{M}^0$ , with time evolution that depends only on  $t_2 - t_1$ .

When  $f_1$  is a state that both  $M^0$  and  $\bar{M}^0$  can decay into, then Eq. (13.37) contains interference terms proportional to  $A_{f_1}\bar{A}_{f_1} \neq 0$  that are not present in Eqs. (13.35, 13.36). Even when  $f_1$  is dominantly produced by  $M^0$  decays rather than  $\bar{M}^0$  decays, or vice versa,  $A_{f_1}\bar{A}_{f_1}$  can be non-zero owing to doubly-CKM-suppressed decays (with amplitudes suppressed by at least two powers of  $\lambda$  relative to the dominant amplitude, in the language of Section 13.3), and these terms should be considered for precision studies of  $CP$  violation in coherent  $V \rightarrow M^0\bar{M}^0$  decays [31]. The correlations in  $V \rightarrow M^0\bar{M}^0$  decays can also be exploited to determine strong phase differences between favored and suppressed decay amplitudes [32].

#### 13.1.4. Classification of $CP$ -violating effects :

We distinguish three types of  $CP$ -violating effects that can occur in the quark sector:

- I.  $CP$  violation in decay is defined by

$$|\bar{A}_f/A_f| \neq 1. \quad (13.39)$$

In charged meson (and all baryon) decays, where mixing effects are absent, this is the only possible source of  $CP$  asymmetries:

$$\mathcal{A}_{f^\pm} \equiv \frac{\Gamma(M^- \rightarrow f^-) - \Gamma(M^+ \rightarrow f^+)}{\Gamma(M^- \rightarrow f^-) + \Gamma(M^+ \rightarrow f^+)} = \frac{|\bar{A}_f/A_f|^2 - 1}{|\bar{A}_f/A_f|^2 + 1}. \quad (13.40)$$

Note that the usual sign convention for  $CP$  asymmetries of hadrons is for the difference between the rate involving the particle that contains a heavy quark and that which contains an antiquark. Hence Eq. (13.40) corresponds to the definition for  $B^\pm$  mesons, but the opposite sign is used for  $D_{(s)}^\pm$  decays.

- II.  $CP$  (and  $T$ ) violation in mixing is defined by

$$|q/p| \neq 1. \quad (13.41)$$

In charged-current semileptonic neutral meson decays  $M, \bar{M} \rightarrow \ell^\pm X$  (taking  $|A_{\ell^+X}| = |\bar{A}_{\ell^-X}|$  and  $A_{\ell^-X} = \bar{A}_{\ell^+X} = 0$ , as is the case in the Standard Model, to lowest order in  $G_F$ , and in most of its reasonable extensions), this is the only source of  $CP$  violation, and can be measured via the asymmetry of “wrong-sign” decays induced by oscillations:

$$\begin{aligned} \mathcal{A}_{\text{SL}}(t) &\equiv \frac{d\Gamma/dt[\bar{M}_{\text{phys}}^0(t) \rightarrow \ell^+ X] - d\Gamma/dt[M_{\text{phys}}^0(t) \rightarrow \ell^- X]}{d\Gamma/dt[\bar{M}_{\text{phys}}^0(t) \rightarrow \ell^+ X] + d\Gamma/dt[M_{\text{phys}}^0(t) \rightarrow \ell^- X]} \\ &= \frac{1 - |q/p|^4}{1 + |q/p|^4}. \end{aligned} \quad (13.42)$$

Note that this asymmetry of time-dependent decay rates is actually time-independent.

- III.  $CP$  violation in interference between a decay without mixing,  $M^0 \rightarrow f$ , and a decay with mixing,  $M^0 \rightarrow \bar{M}^0 \rightarrow f$  (such an effect occurs only in decays to final states that are common to  $M^0$  and  $\bar{M}^0$ , including all  $CP$  eigenstates), is defined by

$$\arg(\lambda_f) + \arg(\lambda_{\bar{f}}) \neq 0, \quad (13.43)$$

with

$$\lambda_f \equiv \frac{q}{p} \frac{\bar{A}_f}{A_f}. \quad (13.44)$$

For final  $CP$  eigenstates,  $f_{CP}$ , the condition Eq. (13.43) simplifies to

$$\text{Im}(\lambda_{f_{CP}}) \neq 0, \quad (13.45)$$

This form of  $CP$  violation can be observed, for example, using the asymmetry of neutral meson decays into  $CP$  eigenstates

$$\mathcal{A}_{f_{CP}}(t) \equiv \frac{d\Gamma/dt[\bar{M}_{\text{phys}}^0(t) \rightarrow f_{CP}] - d\Gamma/dt[M_{\text{phys}}^0(t) \rightarrow f_{CP}]}{d\Gamma/dt[\bar{M}_{\text{phys}}^0(t) \rightarrow f_{CP}] + d\Gamma/dt[M_{\text{phys}}^0(t) \rightarrow f_{CP}]} \quad (13.46)$$

If  $\Delta\Gamma = 0$ , as expected to a good approximation for  $B^0$  mesons, but not for  $K^0$  and  $B_s^0$  mesons, and  $|q/p| = 1$ , then  $\mathcal{A}_{f_{CP}}$  has a particularly simple form (see Eq. (13.91), below). If, in addition, the decay amplitudes fulfill  $|\bar{A}_{f_{CP}}| = |A_{f_{CP}}|$ , the interference between decays with and without mixing is the only source of the asymmetry and  $\mathcal{A}_{f_{CP}}(t) = \text{Im}(\lambda_{f_{CP}}) \sin(x\Gamma t)$ .

Examples of these three types of  $CP$  violation will be given in Sections 13.4, 13.5, and 13.6.

### 13.2. Theoretical Interpretation: General Considerations

Consider the  $M \rightarrow f$  decay amplitude  $A_f$ , and the  $CP$  conjugate process,  $\bar{M} \rightarrow \bar{f}$ , with decay amplitude  $\bar{A}_{\bar{f}}$ . There are two types of phases that may appear in these decay amplitudes. Complex parameters in any Lagrangian term that contributes to the amplitude will appear in complex conjugate form in the  $CP$ -conjugate amplitude. Thus, their phases appear in  $A_f$  and  $\bar{A}_{\bar{f}}$  with opposite signs. In the Standard Model, these phases occur only in the couplings of the  $W^\pm$  bosons, and hence, are often called “weak phases.” The weak phase of any single term is convention-dependent. However, the difference between the weak phases in two different terms in  $A_f$  is convention-independent. A second type of phase can appear in scattering or decay amplitudes, even when the Lagrangian is real. This phase originates from the possible contribution from intermediate on-shell states in the decay process. Since such phases are generated by  $CP$ -invariant interactions, they are the same in  $A_f$  and  $\bar{A}_{\bar{f}}$ . Usually the dominant rescattering is due to strong interactions; hence the designation “strong phases” for the phase shifts so induced. Again, only the relative strong phases between different terms in the amplitude are physically meaningful.

The “weak” and “strong” phases discussed here appear in addition to the spurious  $CP$ -transformation phases of Eq. (13.21). Those spurious phases are due to an arbitrary choice of phase convention, and do not originate from any dynamics or induce any  $CP$  violation. For simplicity, we set them to zero from here on.

It is useful to write each contribution  $a_i$  to  $A_f$  in three parts: its magnitude  $|a_i|$ , its weak phase  $\phi_i$ , and its strong phase  $\delta_i$ . If, for example, there are two such contributions,  $A_f = a_1 + a_2$ , we have

$$\begin{aligned} A_f &= |a_1|e^{i(\delta_1+\phi_1)} + |a_2|e^{i(\delta_2+\phi_2)}, \\ \bar{A}_{\bar{f}} &= |a_1|e^{i(\delta_1-\phi_1)} + |a_2|e^{i(\delta_2-\phi_2)}. \end{aligned} \quad (13.47)$$

Similarly, for neutral mesons, it is useful to write

$$\mathbf{M}_{12} = |\mathbf{M}_{12}|e^{i\phi_M}, \quad \mathbf{\Gamma}_{12} = |\mathbf{\Gamma}_{12}|e^{i\phi_\Gamma}. \quad (13.48)$$

Each of the phases appearing in Eqs. (13.47, 13.48) is convention-dependent, but combinations such as  $\delta_1 - \delta_2$ ,  $\phi_1 - \phi_2$ ,  $\phi_M - \phi_\Gamma$ , and  $\phi_M + \phi_1 - \phi_1$  (where  $\phi_1$  is a weak phase contributing to  $\bar{A}_f$ ) are physical.

It is now straightforward to evaluate the various asymmetries in terms of the theoretical parameters introduced here. We will do so with approximations that are often relevant to the most interesting measured asymmetries.

1. The  $CP$  asymmetry in charged meson and all baryon decays [Eq. (13.40)] is given by

$$\mathcal{A}_f = -\frac{2|a_1a_2|\sin(\delta_2 - \delta_1)\sin(\phi_2 - \phi_1)}{|a_1|^2 + |a_2|^2 + 2|a_1a_2|\cos(\delta_2 - \delta_1)\cos(\phi_2 - \phi_1)}. \quad (13.49)$$

The quantity of most interest to theory is the weak phase difference  $\phi_2 - \phi_1$ . Its extraction from the asymmetry requires, however, that the amplitude ratio  $|a_2/a_1|$  and the strong phase difference  $\delta_2 - \delta_1$  are known. Both quantities depend on non-perturbative hadronic parameters that are difficult to calculate, but in some cases can be obtained from experiment.

2. In the approximation that  $|\mathbf{\Gamma}_{12}/\mathbf{M}_{12}| \ll 1$  (valid for  $B^0$  and  $B_s^0$  mesons), the  $CP$  asymmetry in semileptonic neutral-meson decays [Eq. (13.42)] is given by

$$\mathcal{A}_{\text{SL}} = -\left|\frac{\mathbf{\Gamma}_{12}}{\mathbf{M}_{12}}\right| \sin(\phi_M - \phi_\Gamma). \quad (13.50)$$

The quantity of most interest to theory is the weak phase  $\phi_M - \phi_\Gamma$ . Its extraction from the asymmetry requires, however, that  $|\mathbf{\Gamma}_{12}/\mathbf{M}_{12}|$

is known. This quantity depends on long-distance physics that is difficult to calculate.

3. In the approximations that only a single weak phase contributes to decay,  $A_f = |a_f|e^{i(\delta_f + \phi_f)}$ , and that  $|\mathbf{\Gamma}_{12}/\mathbf{M}_{12}| = 0$ , we obtain  $|\lambda_f| = 1$ , and the  $CP$  asymmetries in decays to a final  $CP$  eigenstate  $f$  [Eq. (13.46)] with eigenvalue  $\eta_f = \pm 1$  are given by

$$\mathcal{A}_{fCP}(t) = \mathcal{I}m(\lambda_f) \sin(\Delta mt) \quad \text{with} \quad \mathcal{I}m(\lambda_f) = \eta_f \sin(\phi_M + 2\phi_f). \quad (13.51)$$

Note that the phase measured is purely a weak phase, and no hadronic parameters are involved in the extraction of its value from  $\mathcal{I}m(\lambda_f)$ .

The discussion above allows us to introduce another classification of  $CP$ -violating effects:

1. *Indirect CP violation* is consistent with taking  $\phi_M \neq 0$  and setting all other  $CP$  violating phases to zero.  $CP$  violation in mixing (type II) belongs to this class.
2. *Direct CP violation* cannot be accounted for by just  $\phi_M \neq 0$ .  $CP$  violation in decay (type I) belongs to this class.

The historical significance of this classification is related to theory. In superweak models [33],  $CP$  violation appears only in diagrams that contribute to  $\mathbf{M}_{12}$ , hence they predict that there is no direct  $CP$  violation. In most models and, in particular, in the Standard Model,  $CP$  violation is both direct and indirect. As concerns type III  $CP$  violation, a single observation of such an effect would be consistent with indirect  $CP$  violation, but observing  $\eta_{f_1} \mathcal{I}m(\lambda_{f_1}) \neq \eta_{f_2} \mathcal{I}m(\lambda_{f_2})$  (for the same decaying meson and two different final  $CP$  eigenstates  $f_1$  and  $f_2$ ) would establish direct  $CP$  violation. The experimental observation of  $\epsilon' \neq 0$ , which was achieved by establishing that  $\mathcal{I}m(\lambda_{\pi^+\pi^-}) \neq \mathcal{I}m(\lambda_{\pi^0\pi^0})$  (see Section 13.4), excluded the superweak scenario.

### 13.3. Theoretical Interpretation: The KM Mechanism

Of all the Standard Model quark parameters, only the Kobayashi-Maskawa (KM) phase is  $CP$ -violating. Having a single source of  $CP$  violation, the Standard Model is very predictive for  $CP$  asymmetries: some vanish, and those that do not are correlated.

To be precise,  $CP$  could be violated also by strong interactions. The experimental upper bound on the electric-dipole moment of the neutron implies, however, that  $\theta_{QCD}$ , the non-perturbative parameter that determines the strength of this type of  $CP$  violation, is tiny, if not zero. (The smallness of  $\theta_{QCD}$  constitutes a theoretical puzzle, known as “the strong  $CP$  problem.”) In particular, it is irrelevant to our discussion of hadron decays.

The charged current interactions (that is, the  $W^\pm$  interactions) for quarks are given by

$$-\mathcal{L}_{W^\pm} = \frac{g}{\sqrt{2}} \overline{u_{Li}} \gamma^\mu (V_{CKM})_{ij} d_{Lj} W_\mu^\pm + \text{h.c.} \quad (13.52)$$

Here  $i, j = 1, 2, 3$  are generation numbers. The Cabibbo-Kobayashi-Maskawa (CKM) mixing matrix for quarks is a  $3 \times 3$  unitary matrix [34]. Ordering the quarks by their masses, *i.e.*,  $(u_1, u_2, u_3) \rightarrow (u, c, t)$  and  $(d_1, d_2, d_3) \rightarrow (d, s, b)$ , the elements of  $V_{CKM}$  are written as follows:

$$V_{CKM} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix}. \quad (13.53)$$

While a general  $3 \times 3$  unitary matrix depends on three real angles and six phases, the freedom to redefine the phases of the quark mass eigenstates can be used to remove five of the phases, leaving a single physical phase, the Kobayashi-Maskawa phase, that is responsible for all  $CP$  violation in the Standard Model.

The fact that one can parametrize  $V_{CKM}$  by three real and only one imaginary physical parameters can be made manifest by choosing an explicit parametrization. The Wolfenstein parametrization [35,36] is particularly useful:

$$V_{CKM} = \begin{pmatrix} 1 - \frac{1}{2}\lambda^2 - \frac{1}{8}\lambda^4 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda + \frac{1}{2}A^2\lambda^5[1 - 2(\rho + i\eta)] & 1 - \frac{1}{2}\lambda^2 - \frac{1}{8}\lambda^4(1 + 4A^2) & A\lambda^2 \\ A\lambda^3[1 - (1 - \frac{1}{2}\lambda^2)(\rho + i\eta)] & -A\lambda^2 + \frac{1}{2}A\lambda^4[1 - 2(\rho + i\eta)] & 1 - \frac{1}{2}A^2\lambda^4 \end{pmatrix}. \quad (13.54)$$

Here  $\lambda \approx 0.23$  (not to be confused with  $\lambda_f$ ), the sine of the Cabibbo angle, plays the role of an expansion parameter, and  $\eta$  represents the  $CP$ -violating phase. Terms of  $\mathcal{O}(\lambda^6)$  have been neglected.

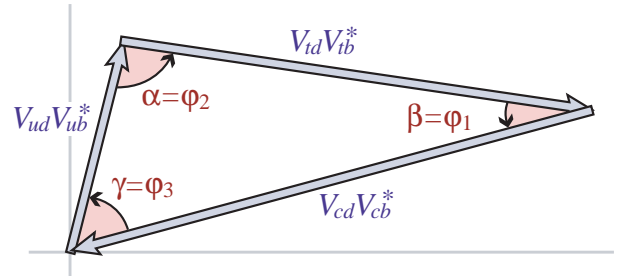
The unitarity of the CKM matrix,  $(VV^\dagger)_{ij} = (V^\dagger V)_{ij} = \delta_{ij}$ , leads to twelve distinct complex relations among the matrix elements. The six relations with  $i \neq j$  can be represented geometrically as triangles in the complex plane. Two of these,

$$\begin{aligned} V_{ud}V_{ub}^* + V_{cd}V_{cb}^* + V_{td}V_{tb}^* &= 0 \\ V_{td}V_{ud}^* + V_{ts}V_{us}^* + V_{tb}V_{ub}^* &= 0, \end{aligned}$$

have terms of equal order,  $\mathcal{O}(A\lambda^3)$ , and so have corresponding triangles whose interior angles are all  $\mathcal{O}(1)$  physical quantities that can be independently measured. The angles of the first triangle (see Fig. 13.1) are given by

$$\begin{aligned} \alpha \equiv \varphi_2 &\equiv \arg\left(-\frac{V_{td}V_{tb}^*}{V_{ud}V_{ub}^*}\right) \simeq \arg\left(-\frac{1 - \rho - i\eta}{\rho + i\eta}\right), \\ \beta \equiv \varphi_1 &\equiv \arg\left(-\frac{V_{cd}V_{cb}^*}{V_{td}V_{tb}^*}\right) \simeq \arg\left(\frac{1}{1 - \rho - i\eta}\right), \\ \gamma \equiv \varphi_3 &\equiv \arg\left(-\frac{V_{ud}V_{ub}^*}{V_{cd}V_{cb}^*}\right) \simeq \arg(\rho + i\eta). \end{aligned} \quad (13.55)$$

The angles of the second triangle are equal to  $(\alpha, \beta, \gamma)$  up to corrections of  $\mathcal{O}(\lambda^2)$ . The notations  $(\alpha, \beta, \gamma)$  and  $(\varphi_1, \varphi_2, \varphi_3)$  are both in common usage but, for convenience, we only use the first convention in the following.



**Figure 13.1:** Graphical representation of the unitarity constraint  $V_{ud}V_{ub}^* + V_{cd}V_{cb}^* + V_{td}V_{tb}^* = 0$  as a triangle in the complex plane.

Another relation that can be represented as a triangle,

$$V_{us}V_{ub}^* + V_{cs}V_{cb}^* + V_{ts}V_{tb}^* = 0, \quad (13.56)$$

and, in particular, its small angle, of  $\mathcal{O}(\lambda^2)$ ,

$$\beta_s \equiv \arg\left(-\frac{V_{ts}V_{tb}^*}{V_{cs}V_{cb}^*}\right), \quad (13.57)$$

is convenient for analyzing  $CP$  violation in the  $B_s^0$  sector.

All unitarity triangles have the same area, commonly denoted by  $J/2$  [37]. If  $CP$  is violated,  $J$  is different from zero and can be taken as the single  $CP$ -violating parameter. In the Wolfenstein parametrization of Eq. (13.54),  $J \simeq \lambda^6 A^2 \eta$ .

### 13.4. Kaons

$CP$  violation was discovered in  $K \rightarrow \pi\pi$  decays in 1964 [1]. The same mode provided the first observation of direct  $CP$  violation [4–6].

The decay amplitudes actually measured in neutral  $K$  decays refer to the mass eigenstates  $K_L$  and  $K_S$ , rather than to the  $K$  and  $\bar{K}$  states referred to in Eq. (13.18). The final  $\pi^+\pi^-$  and  $\pi^0\pi^0$  states are  $CP$ -even. In the  $CP$  conservation limit,  $K_S$  ( $K_L$ ) would be  $CP$ -even (odd), and therefore would (would not) decay to two pions. We define  $CP$ -violating amplitude ratios for two-pion final states,

$$\eta_{00} \equiv \frac{\langle \pi^0\pi^0 | \mathcal{H} | K_L \rangle}{\langle \pi^0\pi^0 | \mathcal{H} | K_S \rangle}, \quad \eta_{+-} \equiv \frac{\langle \pi^+\pi^- | \mathcal{H} | K_L \rangle}{\langle \pi^+\pi^- | \mathcal{H} | K_S \rangle}. \quad (13.58)$$

Another important observable is the asymmetry of time-integrated semileptonic decay rates:

$$\delta_L \equiv \frac{\Gamma(K_L \rightarrow \ell^+\nu_\ell\pi^-) - \Gamma(K_L \rightarrow \ell^-\bar{\nu}_\ell\pi^+)}{\Gamma(K_L \rightarrow \ell^+\nu_\ell\pi^-) + \Gamma(K_L \rightarrow \ell^-\bar{\nu}_\ell\pi^+)}. \quad (13.59)$$

$CP$  violation has been observed as an appearance of  $K_L$  decays to two-pion final states [26],

$$|\eta_{00}| = (2.220 \pm 0.011) \times 10^{-3} \quad |\eta_{+-}| = (2.232 \pm 0.011) \times 10^{-3} \quad (13.60)$$

$$|\eta_{00}/\eta_{+-}| = 0.9950 \pm 0.0007, \quad (13.61)$$

where the phase  $\phi_{ij}$  of the amplitude ratio  $\eta_{ij}$  has been determined both assuming  $CPT$  invariance:

$$\phi_{00} = (43.52 \pm 0.05)^\circ, \quad \phi_{+-} = (43.51 \pm 0.05)^\circ, \quad (13.62)$$

and without assuming  $CPT$  invariance:

$$\phi_{00} = (43.7 \pm 0.6)^\circ, \quad \phi_{+-} = (43.4 \pm 0.5)^\circ. \quad (13.63)$$

$CP$  violation has also been observed in semileptonic  $K_L$  decays [26]

$$\delta_L = (3.32 \pm 0.06) \times 10^{-3}, \quad (13.64)$$

where  $\delta_L$  is a weighted average of muon and electron measurements, as well as in  $K_L$  decays to  $\pi^+\pi^-\gamma$  and  $\pi^+\pi^-e^+e^-$  [26].  $CP$  violation in  $K \rightarrow 3\pi$  decays has not yet been observed [26,38].

Historically,  $CP$  violation in neutral  $K$  decays has been described in terms of the complex parameters  $\epsilon$  and  $\epsilon'$ . The observables  $\eta_{00}$ ,  $\eta_{+-}$ , and  $\delta_L$  are related to these parameters, and to those of Section 13.1, by

$$\begin{aligned} \eta_{00} &= \frac{1 - \lambda_{\pi^0\pi^0}}{1 + \lambda_{\pi^0\pi^0}} = \epsilon - 2\epsilon', \\ \eta_{+-} &= \frac{1 - \lambda_{\pi^+\pi^-}}{1 + \lambda_{\pi^+\pi^-}} = \epsilon + \epsilon', \\ \delta_L &= \frac{1 - |q/p|^2}{1 + |q/p|^2} = \frac{2\mathcal{R}e(\epsilon)}{1 + |\epsilon|^2}, \end{aligned} \quad (13.65)$$

where, in the last line, we have assumed that  $|A_{\ell^+\nu_\ell\pi^-}| = |\bar{A}_{\ell^-\bar{\nu}_\ell\pi^+}|$  and  $|A_{\ell^-\bar{\nu}_\ell\pi^+}| = |\bar{A}_{\ell^+\nu_\ell\pi^-}| = 0$ . (The convention-dependent parameter  $\bar{\epsilon} \equiv (1 - q/p)/(1 + q/p)$ , sometimes used in the literature, is, in general, different from  $\epsilon$  but yields a similar expression,  $\delta_L = 2\mathcal{R}e(\bar{\epsilon})/(1 + |\bar{\epsilon}|^2)$ .) A fit to the  $K \rightarrow \pi\pi$  data yields [26]

$$\begin{aligned} |\epsilon| &= (2.228 \pm 0.011) \times 10^{-3}, \\ \mathcal{R}e(\epsilon'/\epsilon) &= (1.66 \pm 0.23) \times 10^{-3}. \end{aligned} \quad (13.66)$$

In discussing two-pion final states, it is useful to express the amplitudes  $A_{\pi^0\pi^0}$  and  $A_{\pi^+\pi^-}$  in terms of their isospin components via

$$\begin{aligned} A_{\pi^0\pi^0} &= \sqrt{\frac{1}{3}} |A_0| e^{i(\delta_0+\phi_0)} - \sqrt{\frac{2}{3}} |A_2| e^{i(\delta_2+\phi_2)}, \\ A_{\pi^+\pi^-} &= \sqrt{\frac{2}{3}} |A_0| e^{i(\delta_0+\phi_0)} + \sqrt{\frac{1}{3}} |A_2| e^{i(\delta_2+\phi_2)}, \end{aligned} \quad (13.67)$$

where we parameterize the amplitude  $A_I(\bar{A}_I)$  for  $K^0(\bar{K}^0)$  decay into two pions with total isospin  $I = 0$  or  $2$  as

$$\begin{aligned} A_I &\equiv \langle (\pi\pi)_I | \mathcal{H} | K^0 \rangle = |A_I| e^{i(\delta_I+\phi_I)}, \\ \bar{A}_I &\equiv \langle (\pi\pi)_I | \mathcal{H} | \bar{K}^0 \rangle = |A_I| e^{i(\delta_I-\phi_I)}. \end{aligned} \quad (13.68)$$

The smallness of  $|\eta_{00}|$  and  $|\eta_{+-}|$  allows us to approximate

$$\epsilon \simeq \frac{1}{2}(1 - \lambda_{(\pi\pi)_{I=0}}), \quad \epsilon' \simeq \frac{1}{6}(\lambda_{\pi^0\pi^0} - \lambda_{\pi^+\pi^-}). \quad (13.69)$$

The parameter  $\epsilon$  represents indirect  $CP$  violation, while  $\epsilon'$  parameterizes direct  $CP$  violation:  $\mathcal{R}e(\epsilon')$  measures  $CP$  violation in decay (type I),  $\mathcal{R}e(\epsilon)$  measures  $CP$  violation in mixing (type II), and  $\mathcal{I}m(\epsilon)$  and  $\mathcal{I}m(\epsilon')$  measure the interference between decays with and without mixing (type III).

The following expressions for  $\epsilon$  and  $\epsilon'$  are useful for theoretical evaluations:

$$\epsilon \simeq \frac{e^{i\pi/4}}{\sqrt{2}} \frac{\mathcal{I}m(\mathbf{M}_{12})}{\Delta m}, \quad \epsilon' = \frac{i}{\sqrt{2}} \left| \frac{A_2}{A_0} \right| e^{i(\delta_2-\delta_0)} \sin(\phi_2 - \phi_0). \quad (13.70)$$

The expression for  $\epsilon$  is only valid in a phase convention where  $\phi_2 = 0$ , corresponding to a real  $V_{ud}V_{us}^*$ , and in the approximation that also  $\phi_0 = 0$ . The phase of  $\epsilon$ ,  $\arg(\epsilon) \approx \arctan(-2\Delta m/\Delta\Gamma)$ , is independent of the electroweak model and is experimentally determined to be about  $\pi/4$ . The calculation of  $\epsilon$  benefits from the fact that  $\mathcal{I}m(\mathbf{M}_{12})$  is dominated by short distance physics. Consequently, the main sources of uncertainty in theoretical interpretations of  $\epsilon$  are the values of matrix elements, such as  $\langle K^0 | (\bar{s}d)_{V-A} (\bar{s}d)_{V-A} | \bar{K}^0 \rangle$ . The expression for  $\epsilon'$  is valid to first order in  $|A_2/A_0| \sim 1/20$ . The phase of  $\epsilon'$  is experimentally determined,  $\pi/2 + \delta_2 - \delta_0 \approx \pi/4$ , and is independent of the electroweak model. Note that, accidentally,  $\epsilon'/\epsilon$  is real to a good approximation. Determination of weak phase information from the measurement of  $\mathcal{R}e(\epsilon'/\epsilon)$  given in Eq. (13.66) has until now been precluded by uncertainties in the hadronic parameters, but recent advances in lattice QCD calculations [39,40] suggest that it may become possible [41].

A future measurement of much interest is that of  $CP$  violation in the rare  $K \rightarrow \pi\nu\bar{\nu}$  decays. The signal for  $CP$  violation is simply observing the  $K_L \rightarrow \pi^0\nu\bar{\nu}$  decay. The effect here is that of interference between decays with and without mixing (type III) [42]:

$$\frac{\Gamma(K_L \rightarrow \pi^0\nu\bar{\nu})}{\Gamma(K^+ \rightarrow \pi^+\nu\bar{\nu})} = \frac{1}{2} [1 + |\lambda_{\pi\nu\bar{\nu}}|^2 - 2\mathcal{R}e(\lambda_{\pi\nu\bar{\nu}})] \simeq 1 - \mathcal{R}e(\lambda_{\pi\nu\bar{\nu}}), \quad (13.71)$$

where in the last equation we neglect  $CP$  violation in decay and in mixing (expected, model-independently, to be of order  $10^{-5}$  and  $10^{-3}$ , respectively). Such a measurement is experimentally very challenging but would be theoretically very rewarding [43]. Similar to the  $CP$  asymmetry in  $B^0 \rightarrow J/\psi K_S$ , the  $CP$  violation in  $K \rightarrow \pi\nu\bar{\nu}$  decay is predicted to be large (that is, the ratio in Eq. (13.71) is neither CKM-nor loop-suppressed) and can be very cleanly interpreted.

Within the Standard Model, the  $K_L \rightarrow \pi^0\nu\bar{\nu}$  decay is dominated by an intermediate top quark contribution and, consequently, can be interpreted in terms of CKM parameters [44]. (For the charged mode,  $K^+ \rightarrow \pi^+\nu\bar{\nu}$ , the contribution from an intermediate charm quark is not negligible, and constitutes a source of hadronic uncertainty.) In particular,  $\mathcal{B}(K_L \rightarrow \pi^0\nu\bar{\nu})$  provides a theoretically clean way to determine the Wolfenstein parameter  $\eta$  [45]:

$$\mathcal{B}(K_L \rightarrow \pi^0\nu\bar{\nu}) = \kappa_L [X(m_t^2/m_W^2)]^2 A^4 \eta^2, \quad (13.72)$$

where the hadronic parameter  $\kappa_L \sim 2 \times 10^{-10}$  incorporates the value of the four-fermion matrix element which is deduced, using isospin relations, from  $\mathcal{B}(K^+ \rightarrow \pi^0 e^+ \nu_e)$ , and  $X(m_t^2/m_W^2)$  is a known function of the top mass. An explicit calculation gives  $\mathcal{B}(K_L \rightarrow \pi^0\nu\bar{\nu}) = (2.4 \pm 0.4) \times 10^{-11}$  [46]. The currently tightest experimental limit is  $\mathcal{B}(K_L \rightarrow \pi^0\nu\bar{\nu}) < 2.6 \times 10^{-8}$  [47], which does not yet reach the bound  $\mathcal{B}(K_L \rightarrow \pi^0\nu\bar{\nu}) < 4.4 \times \mathcal{B}(K^+ \rightarrow \pi^+\nu\bar{\nu})$  [42]. Significant further progress is anticipated from experiments searching for  $K \rightarrow \pi\nu\bar{\nu}$  decays in the next few years [48,49].

### 13.5. Charm

The existence of  $D^0\text{--}\bar{D}^0$  mixing has been established in recent years [50–53]. The experimental constraints read [28,54]  $x \equiv \Delta m/\Gamma = (0.37 \pm 0.16) \times 10^{-2}$  and  $y \equiv \Delta\Gamma/(2\Gamma) = (0.66^{+0.07}_{-0.10}) \times 10^{-2}$ . Thus, the data clearly show that  $y \neq 0$ , but improved measurements are needed to be sure of the size of  $x$ . Long-distance contributions make it difficult to calculate Standard Model predictions for the  $D^0\text{--}\bar{D}^0$  mixing parameters. Therefore, the goal of the search for  $D^0\text{--}\bar{D}^0$  mixing is not to constrain the CKM parameters, but rather to probe new physics. Here *CP* violation plays an important role. Within the Standard Model, the *CP*-violating effects are predicted to be small, since the mixing and the relevant decays are described, to an excellent approximation, by the physics of the first two generations only. The expectation is that the Standard Model size of *CP* violation in  $D$  decays is  $\mathcal{O}(10^{-3})$  or less, but theoretical work is ongoing to understand whether QCD effects can significantly enhance it. At present, the most sensitive searches involve the  $D^0 \rightarrow K^+K^-$ ,  $D^0 \rightarrow \pi^+\pi^-$  and  $D^0 \rightarrow K^\pm\pi^\mp$  modes.

The neutral  $D$  mesons decay via a singly-Cabibbo-suppressed transition to the *CP* eigenstates  $K^+K^-$  and  $\pi^+\pi^-$ . These decays are dominated by Standard-Model tree diagrams. Thus, we can write, for  $f = K^+K^-$  or  $\pi^+\pi^-$ ,

$$\begin{aligned} A_f &= A_f^T e^{+i\phi_f^T} \left[ 1 + r_f e^{i(\delta_f + \phi_f)} \right], \\ \bar{A}_f &= A_f^T e^{-i\phi_f^T} \left[ 1 + r_f e^{i(\delta_f - \phi_f)} \right], \end{aligned} \quad (13.73)$$

where  $A_f^T e^{\pm i\phi_f^T}$  is the Standard Model tree-level contribution,  $\phi_f^T$  and  $\phi_f$  are weak, *CP* violating phases,  $\delta_f$  is a strong phase difference, and  $r_f$  is the ratio between a subleading ( $r_f \ll 1$ ) contribution with a weak phase different from  $\phi_f^T$  and the Standard Model tree-level contribution. Neglecting  $r_f$ ,  $\lambda_f$  is universal, and we can define an observable phase  $\phi_D$  via

$$\lambda_f \equiv -|q/p| e^{i\phi_D}. \quad (13.74)$$

(In the limit of *CP* conservation, choosing  $\phi_D = 0$  is equivalent to defining the mass eigenstates by their *CP* eigenvalue:  $|D_\pm\rangle = p|D^0\rangle \pm q|\bar{D}^0\rangle$ , with  $D_-$  ( $D_+$ ) being the *CP*-odd (*CP*-even) state; that is, the state that does not (does) decay into  $K^+K^-$ .)

We define the time integrated *CP* asymmetry for a final *CP* eigenstate  $f$  as follows:

$$a_f \equiv \frac{\int_0^\infty \Gamma(D_{\text{phys}}^0(t) \rightarrow f) dt - \int_0^\infty \Gamma(\bar{D}_{\text{phys}}^0(t) \rightarrow f) dt}{\int_0^\infty \Gamma(D_{\text{phys}}^0(t) \rightarrow f) dt + \int_0^\infty \Gamma(\bar{D}_{\text{phys}}^0(t) \rightarrow f) dt}. \quad (13.75)$$

(This expression corresponds to the  $D$  meson being tagged at production, hence the integration goes from 0 to  $+\infty$ ; measurements are also possible with  $\psi(3770) \rightarrow D^0\bar{D}^0$ , in which case the integration goes from  $-\infty$  to  $+\infty$  giving slightly different results; see the discussion in Section 13.1.3.) We take  $x, y, r_f \ll 1$  and expand to leading order in these parameters. We can then separate the contribution to  $a_f$  into three parts [55],

$$a_f = a_f^d + a_f^m + a_f^i, \quad (13.76)$$

with the following underlying mechanisms:

1.  $a_f^d$  signals *CP* violation in decay (similar to Eq. (13.40)):

$$a_f^d = 2r_f \sin \phi_f \sin \delta_f. \quad (13.77)$$

2.  $a_f^m$  signals *CP* violation in mixing (similar to Eq. (13.50)). With our approximations, it is universal:

$$a^m = -\frac{y}{2} \left( \left| \frac{q}{p} \right| - \left| \frac{p}{q} \right| \right) \cos \phi_D. \quad (13.78)$$

3.  $a_f^i$  signals *CP* violation in the interference of mixing and decay (similar to Eq. (13.51)). With our approximations, it is universal:

$$a^i = \frac{x}{2} \left( \left| \frac{q}{p} \right| + \left| \frac{p}{q} \right| \right) \sin \phi_D. \quad (13.79)$$

One can isolate the effects of direct *CP* violation by taking the difference between the *CP* asymmetries in the  $K^+K^-$  and  $\pi^+\pi^-$  modes:

$$\Delta a_{CP} \equiv a_{K^+K^-} - a_{\pi^+\pi^-} = a_{K^+K^-}^d - a_{\pi^+\pi^-}^d, \quad (13.80)$$

where we neglected a residual, experiment-dependent, contribution from indirect *CP* violation due to the fact that there may be a decay time-dependent acceptance function that can be different for the  $K^+K^-$  and  $\pi^+\pi^-$  channels. Recent evidence for such direct *CP* violation [56] has become less significant when including more data, with the current average giving [28]:

$$a_{K^+K^-}^d - a_{\pi^+\pi^-}^d = (2.6 \pm 1.0) \times 10^{-3}. \quad (13.81)$$

One can also isolate the effects of indirect *CP* violation in the following way. Consider the time-dependent decay rates in Eq. (13.35) and Eq. (13.36). The mixing processes modify the time dependence from a pure exponential. However, given the small values of  $x$  and  $y$ , the time dependences can be recast, to a good approximation, into purely exponential form, but with modified decay-rate parameters [57,58] (given here for the  $K^+K^-$  final state):

$$\begin{aligned} \Gamma_{D^0 \rightarrow K^+K^-} &= \Gamma \times [1 + |q/p| (y \cos \phi_D - x \sin \phi_D)], \\ \Gamma_{\bar{D}^0 \rightarrow K^+K^-} &= \Gamma \times [1 + |p/q| (y \cos \phi_D + x \sin \phi_D)]. \end{aligned} \quad (13.82)$$

One can define *CP*-conserving and *CP*-violating combinations of these two observables (normalized to the true width  $\Gamma$ ):

$$\begin{aligned} y_{CP} &\equiv \frac{\Gamma_{\bar{D}^0 \rightarrow K^+K^-} + \Gamma_{D^0 \rightarrow K^+K^-}}{2\Gamma} - 1 \\ &= (y/2) (|q/p| + |p/q|) \cos \phi_D - (x/2) (|q/p| - |p/q|) \sin \phi_D, \\ A_\Gamma &\equiv \frac{\Gamma_{D^0 \rightarrow K^+K^-} - \Gamma_{\bar{D}^0 \rightarrow K^+K^-}}{2\Gamma} \\ &= -(a^m + a^i). \end{aligned} \quad (13.83)$$

In the limit of *CP* conservation (and, in particular, within the Standard Model),  $y_{CP} = (\Gamma_+ - \Gamma_-)/2\Gamma = y$  (where  $\Gamma_+(\Gamma_-)$  is the decay width of the *CP*-even (-odd) mass eigenstate) and  $A_\Gamma = 0$ . Indeed, present measurements imply that *CP* violation is small [28],

$$\begin{aligned} y_{CP} &= (+0.84 \pm 0.16) \times 10^{-2}, \\ A_\Gamma &= (-0.06 \pm 0.04) \times 10^{-2}. \end{aligned}$$

The  $K^\pm\pi^\mp$  states are not *CP* eigenstates, but they are still common final states for  $D^0$  and  $\bar{D}^0$  decays. Since  $D^0(\bar{D}^0) \rightarrow K^-\pi^+$  is a Cabibbo-favored (doubly-Cabibbo-suppressed) process, these processes are particularly sensitive to  $x$  and/or  $y = \mathcal{O}(\lambda^2)$ . Taking into account that  $|\lambda_{K^-\pi^+}|, |\lambda_{K^+\pi^-}| \ll 1$  and  $x, y \ll 1$ , assuming that there is no direct *CP* violation (these are Standard Model tree-level decays dominated by a single weak phase, and there is no contribution from penguin-like and chromomagnetic operators), and expanding the time-dependent rates for  $xt, yt \lesssim \Gamma^{-1}$ , one obtains

$$\begin{aligned} \Gamma[D_{\text{phys}}^0(t) \rightarrow K^+\pi^-] &= e^{-\Gamma t} |\bar{A}_{K^-\pi^+}|^2 \\ &\times \left[ r_d^2 + r_d \left| \frac{q}{p} \right| (y' \cos \phi_D - x' \sin \phi_D) \Gamma t + \left| \frac{q}{p} \right|^2 \frac{y^2 + x^2}{4} (\Gamma t)^2 \right], \\ \Gamma[\bar{D}_{\text{phys}}^0(t) \rightarrow K^-\pi^+] &= e^{-\Gamma t} |\bar{A}_{K^-\pi^+}|^2 \\ &\times \left[ r_d^2 + r_d \left| \frac{p}{q} \right| (y' \cos \phi_D + x' \sin \phi_D) \Gamma t + \left| \frac{p}{q} \right|^2 \frac{y^2 + x^2}{4} (\Gamma t)^2 \right], \end{aligned}$$

where

$$\begin{aligned} y' &\equiv y \cos \delta - x \sin \delta, \\ x' &\equiv x \cos \delta + y \sin \delta. \end{aligned} \quad (13.85)$$

The weak phase  $\phi_D$  is the same as that of Eq. (13.74) (a consequence of neglecting direct *CP* violation) and  $r_d = \mathcal{O}(\tan^2 \theta_c)$  is the amplitude ratio,  $r_d = |\bar{A}_{K^-\pi^+}/A_{K^-\pi^+}| = |A_{K^+\pi^-}/\bar{A}_{K^+\pi^-}|$ , that

is,  $\lambda_{K^-\pi^+} = r_d |q/p| e^{-i(\delta-\phi_D)}$  and  $\lambda_{K^+\pi^-}^{-1} = r_d |p/q| e^{-i(\delta+\phi_D)}$ . The parameter  $\delta$  is a strong-phase difference for these processes, that can be obtained from measurements of quantum correlated  $\psi(3770) \rightarrow D^0 \bar{D}^0$  decays [59,60]. By fitting to the six coefficients of the various time-dependences, one can determine  $r_d$ ,  $|q/p|$ ,  $(x^2 + y^2)$ ,  $y' \cos \phi_D$ , and  $x' \sin \phi_D$ . In particular, finding *CP* violation ( $|q/p| \neq 1$  and/or  $\sin \phi_D \neq 0$ ) at a level much higher than  $10^{-3}$  would constitute evidence for new physics. The most stringent constraints to date on *CP* violation in charm mixing have been obtained with this method [61].

A fit to all data [28], including also results from time-dependent analyses of  $D^0 \rightarrow K_S \pi^+ \pi^-$  decays, from which  $x$ ,  $y$ ,  $|q/p|$  and  $\phi_D$  can be determined directly, yields no evidence for indirect *CP* violation:

$$1 - |q/p| = +0.09_{-0.12}^{+0.08},$$

$$\phi_D = \left(-9_{-10}^{+12}\right)^\circ.$$

With the additional assumption of no direct *CP* violation in doubly-Cabibbo-suppressed *D* decays [62–64], tighter constraints are obtained:

$$1 - |q/p| = -0.002 \pm 0.014,$$

$$\phi_D = (-0.1 \pm 0.6)^\circ.$$

More details on various theoretical and experimental aspects of  $D^0 - \bar{D}^0$  mixing can be found in Ref. [30].

Searches for *CP* violation in charged  $D_{(s)}$  decays have been performed in many modes. Searches in decays mediated by Cabibbo-suppressed amplitudes are particularly interesting, since in other channels effects are likely to be too small to be observable in current experiments. Examples of relevant two-body modes are  $D^+ \rightarrow \pi^+ \pi^0$ ,  $K_S K^+$ ,  $\phi \pi^+$  and  $D_s^+ \rightarrow K^+ \pi^0$ ,  $K_S \pi^+$ ,  $\phi K^+$ . The most precise results are  $\mathcal{A}_{D^+ \rightarrow K_S K^+} = -0.0003 \pm 0.0017$  and  $\mathcal{A}_{D_s^+ \rightarrow K_S \pi^+} = +0.0063 \pm 0.0047$  [28]. The precision of experiments is now sufficient that the effect from *CP* violation in the neutral kaon system can be seen in  $D^+ \rightarrow K_S \pi^+$  decays [65,66].

Three-body final states provide additional possibilities to search for *CP* violation, since effects may vary over the phase-space. A number of methods have been proposed to exploit this feature and search for *CP* violation in ways that do not require modelling of the decay distribution [67–69]. Such methods are useful for analysis of charm decays since they are less sensitive to biases from production asymmetries, and are well suited to address the issue of whether or not *CP* violation effects are present. The results of all searches to date have been null – no significant *CP* violation effect has yet been observed in  $D_{(s)}^+$  decays.

## 13.6. Beauty

### 13.6.1. *CP violation in mixing of $B^0$ and $B_s^0$ mesons*

The upper bound on the *CP* asymmetry in semileptonic *B* decays [27] implies that *CP* violation in  $B^0 - \bar{B}^0$  mixing is a small effect (we use  $\mathcal{A}_{\text{SL}}/2 \approx 1 - |q/p|$ , see Eq. (13.42)):

$$\mathcal{A}_{\text{SL}}^d = (-1.5 \pm 1.7) \times 10^{-3} \implies |q/p| = 1.0007 \pm 0.0009. \quad (13.86)$$

The Standard Model prediction is

$$\mathcal{A}_{\text{SL}}^d = \mathcal{O} \left[ (m_c^2/m_t^2) \sin \beta \right] \lesssim 0.001. \quad (13.87)$$

An explicit calculation gives  $(-4.1 \pm 0.6) \times 10^{-4}$  [70].

The experimental constraint on *CP* violation in  $B_s^0 - \bar{B}_s^0$  mixing is somewhat weaker than that in the  $B^0 - \bar{B}^0$  system [27]

$$\mathcal{A}_{\text{SL}}^s = (-7.5 \pm 4.1) \times 10^{-3} \implies |q/p| = 1.0038 \pm 0.0021. \quad (13.88)$$

The Standard Model prediction is  $\mathcal{A}_{\text{SL}}^s = \mathcal{O} \left[ (m_c^2/m_t^2) \sin \beta_s \right] \lesssim 10^{-4}$ , with an explicit calculation giving  $(1.9 \pm 0.3) \times 10^{-5}$  [70]. The tension between the measurement and the prediction originates from a result

from D0 for the inclusive same-sign dimuon asymmetry that deviates from the Standard Model prediction by  $3.6\sigma$  [71]. As yet, this has not been confirmed by independent studies.

In models where  $\Gamma_{12}/\mathbf{M}_{12}$  is approximately real, such as the Standard Model, an upper bound on  $\Delta\Gamma/\Delta m \approx \text{Re}(\Gamma_{12}/\mathbf{M}_{12})$  provides yet another upper bound on the deviation of  $|q/p|$  from one. This constraint does not hold if  $\Gamma_{12}/\mathbf{M}_{12}$  is approximately imaginary. (An alternative parameterization uses  $q/p = (1 - \bar{\epsilon}_B)/(1 + \bar{\epsilon}_B)$ , leading to  $\mathcal{A}_{\text{SL}} \simeq 4\text{Re}(\bar{\epsilon}_B)$ .)

### 13.6.2. *CP violation in interference of $B^0$ decays with and without mixing*

The small deviation (less than one percent) of  $|q/p|$  from 1 implies that, at the present level of experimental precision, *CP* violation in  $B^0$  mixing is a negligible effect. Thus, for the purpose of analyzing *CP* asymmetries in hadronic  $B^0$  decays, we can use

$$\lambda_f = e^{-i\phi_{M(B^0)}} (\bar{A}_f/A_f), \quad (13.89)$$

where  $\phi_{M(B^0)}$  refers to the phase of  $\mathbf{M}_{12}$  appearing in Eq. (13.48) that is appropriate for  $B^0 - \bar{B}^0$  oscillations. Within the Standard Model, the corresponding phase factor is given by

$$e^{-i\phi_{M(B^0)}} = (V_{tb}^* V_{td}) / (V_{tb} V_{td}^*). \quad (13.90)$$

The class of *CP* violation effects in interference between mixing and decay is studied with final states that are common to  $B^0$  and  $\bar{B}^0$  decays [72,73]. It is convenient to rewrite Eq. (13.46) for  $B^0$  decays as [74–76]

$$\mathcal{A}_f(t) = S_f \sin(\Delta m t) - C_f \cos(\Delta m t),$$

$$S_f \equiv \frac{2\text{Im}(\lambda_f)}{1 + |\lambda_f|^2}, \quad C_f \equiv \frac{1 - |\lambda_f|^2}{1 + |\lambda_f|^2}, \quad (13.91)$$

where we assume that  $\Delta\Gamma = 0$  and  $|q/p| = 1$ . An alternative notation in use is  $A_f \equiv -C_f$  – this  $A_f$  should not be confused with the  $A_f$  of Eq. (13.18), but in the limit that  $|q/p| = 1$  is equivalent with the  $\mathcal{A}_f$  of Eq. (13.40).

A large class of interesting processes proceed via quark transitions of the form  $\bar{b} \rightarrow \bar{q} q q'$  with  $q' = s$  or  $d$ . For  $q = c$  or  $u$ , there are contributions from both tree ( $t$ ) and penguin ( $p^{qu}$ , where  $q_u = u, c, t$  is the quark in the loop) diagrams (see Fig. 13.2) which carry different weak phases:

$$A_f = \left( V_{qb}^* V_{qq'} \right) t_f + \sum_{qu=u,c,t} \left( V_{qb}^* V_{quq'} \right) p_f^{qu}. \quad (13.92)$$

(The distinction between tree and penguin contributions is a heuristic one; the separation by the operator that enters is more precise. A detailed discussion of the more complete operator product approach, which also includes higher order QCD corrections, can be found in Ref. [77] for example.) Using CKM unitarity, these decay amplitudes can always be written in terms of just two CKM combinations. For example, for  $f = \pi\pi$ , which proceeds via a  $\bar{b} \rightarrow \bar{u} u \bar{d}$  transition, we can write

$$A_{\pi\pi} = (V_{ub}^* V_{ud}) T_{\pi\pi} + (V_{tb}^* V_{td}) P_{\pi\pi}^t, \quad (13.93)$$

where  $T_{\pi\pi} = t_{\pi\pi} + p_{\pi\pi}^u - p_{\pi\pi}^c$  and  $P_{\pi\pi}^t = p_{\pi\pi}^t - p_{\pi\pi}^c$ . *CP*-violating phases in Eq. (13.93) appear only in the CKM elements, so that

$$\frac{\bar{A}_{\pi\pi}}{A_{\pi\pi}} = \frac{(V_{ub} V_{ud}^*) T_{\pi\pi} + (V_{tb} V_{td}^*) P_{\pi\pi}^t}{(V_{ub}^* V_{ud}) T_{\pi\pi} + (V_{tb}^* V_{td}) P_{\pi\pi}^t}. \quad (13.94)$$

For  $f = J/\psi K$ , which proceeds via a  $\bar{b} \rightarrow \bar{c} c \bar{s}$  transition, we can write

$$A_{\psi K} = (V_{cb}^* V_{cs}) T_{\psi K} + (V_{ub}^* V_{us}) P_{\psi K}^u, \quad (13.95)$$

where  $T_{\psi K} = t_{\psi K} + p_{\psi K}^c - p_{\psi K}^t$  and  $P_{\psi K}^u = p_{\psi K}^u - p_{\psi K}^t$ . A subtlety arises in this decay that is related to the fact that  $B^0$  decays into a final  $J/\psi K^0$  state while  $\bar{B}^0$  decays into a final  $J/\psi \bar{K}^0$  state. A common final state, *e.g.*,  $J/\psi K_S$ , is reached only via  $K^0 - \bar{K}^0$  mixing.



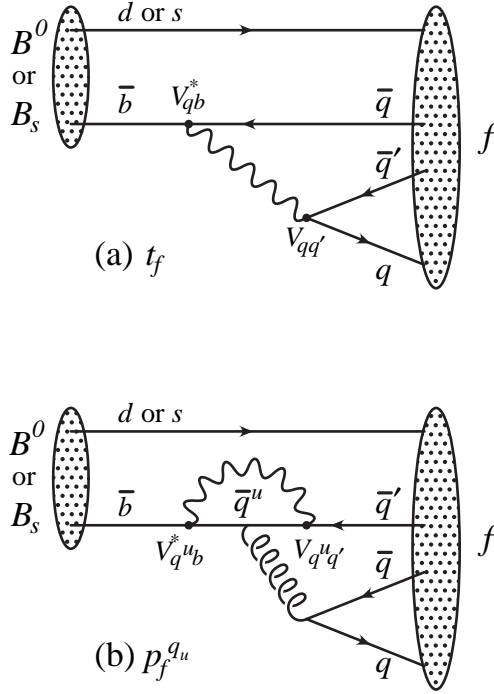
Consequently, the phase factor (defined in Eq. (13.48)) corresponding to neutral  $K$  mixing,  $e^{-i\phi_M(K)} = (V_{cd}^* V_{cs}) / (V_{cd} V_{cs}^*)$ , plays a role:

$$\frac{\bar{A}_{\psi K_S}}{A_{\psi K_S}} = -\frac{(V_{cb}^* V_{cs}) T_{\psi K} + (V_{ub} V_{us}^*) P_{\psi K}^u}{(V_{cb}^* V_{cs}) T_{\psi K} + (V_{ub}^* V_{us}) P_{\psi K}^u} \times \frac{V_{cd}^* V_{cs}}{V_{cd} V_{cs}^*}. \quad (13.96)$$

For  $q = s$  or  $d$ , there are only penguin contributions to  $A_f$ , that is,  $t_f = 0$  in Eq. (13.92). (The tree  $\bar{b} \rightarrow \bar{u}u\bar{q}'$  transition followed by  $\bar{u}u \rightarrow \bar{q}q$  rescattering is included below in the  $P^u$  terms.) Again, CKM unitarity allows us to write  $A_f$  in terms of two CKM combinations. For example, for  $f = \phi K_S$ , which proceeds via a  $\bar{b} \rightarrow \bar{s}s\bar{s}$  transition, we can write

$$\frac{\bar{A}_{\phi K_S}}{A_{\phi K_S}} = -\frac{(V_{cb}^* V_{cs}) P_{\phi K}^c + (V_{ub} V_{us}^*) P_{\phi K}^u}{(V_{cb}^* V_{cs}) P_{\phi K}^c + (V_{ub}^* V_{us}) P_{\phi K}^u} \times \frac{V_{cd}^* V_{cs}}{V_{cd} V_{cs}^*}, \quad (13.97)$$

where  $P_{\phi K}^c = p_{\phi K}^c - p_{\phi K}^t$  and  $P_{\phi K}^u = p_{\phi K}^u - p_{\phi K}^t$ .



**Figure 13.2:** Feynman diagrams for (a) tree and (b) penguin amplitudes contributing to  $B^0 \rightarrow f$  or  $B_s^0 \rightarrow f$  via a  $\bar{b} \rightarrow \bar{q}q'$  quark-level process.

Since in general the amplitude  $A_f$  involves two different weak phases, the corresponding decays can exhibit both  $CP$  violation in the interference of decays with and without mixing,  $S_f \neq 0$ , and  $CP$  violation in decay,  $C_f \neq 0$ . (At the present level of experimental precision, the contribution to  $C_f$  from  $CP$  violation in mixing is negligible, see Eq. (13.86).) If the contribution from a second weak phase is suppressed, then the interpretation of  $S_f$  in terms of Lagrangian  $CP$ -violating parameters is clean, while  $C_f$  is small. If such a second contribution is not suppressed,  $S_f$  depends on hadronic parameters and, if the relevant strong phase difference is large,  $C_f$  is large.

A summary of  $\bar{b} \rightarrow \bar{q}q'$  modes with  $q' = s$  or  $d$  is given in Table 13.1. The  $\bar{b} \rightarrow \bar{d}d\bar{q}$  transitions lead to final states that are similar to those from  $\bar{b} \rightarrow \bar{u}u\bar{q}$  transitions and have similar phase dependence. Final states that consist of two vector mesons ( $\psi\phi$  and  $\phi\phi$ ) are not  $CP$  eigenstates, and angular analysis is needed to separate the  $CP$ -even from the  $CP$ -odd contributions.

**Table 13.1:** Summary of  $\bar{b} \rightarrow \bar{q}q'$  modes with  $q' = s$  or  $d$ . The second and third columns give examples of hadronic final states (usually those which are experimentally most convenient to study). The fourth column gives the CKM dependence of the amplitude  $A_f$ , using the notation of Eqs. (13.93, 13.95, 13.97), with the dominant term first and the subdominant second. The suppression factor of the second term compared to the first is given in the last column. “Loop” refers to a penguin versus tree-suppression factor (it is mode-dependent and roughly  $\mathcal{O}(0.2 - 0.3)$ ) and  $\lambda \simeq 0.23$  is the expansion parameter of Eq. (13.54).

$\bar{b} \rightarrow \bar{q}q'$	$B^0 \rightarrow f$	$B_s^0 \rightarrow f$	CKM dependence of $A_f$	Suppression
$\bar{b} \rightarrow \bar{c}c\bar{s}$	$\psi K_S$	$\psi\phi$	$(V_{cb}^* V_{cs})T + (V_{ub}^* V_{us})P^u$	$\text{loop} \times \lambda^2$
$\bar{b} \rightarrow \bar{s}s\bar{s}$	$\phi K_S$	$\phi\phi$	$(V_{cb}^* V_{cs})P^c + (V_{ub}^* V_{us})P^u$	$\lambda^2$
$\bar{b} \rightarrow \bar{u}u\bar{s}$	$\pi^0 K_S$	$K^+ K^-$	$(V_{cb}^* V_{cs})P^c + (V_{ub}^* V_{us})T$	$\lambda^2/\text{loop}$
$\bar{b} \rightarrow \bar{c}c\bar{d}$	$D^+ D^-$	$\psi K_S$	$(V_{cb}^* V_{cd})T + (V_{ub}^* V_{ud})P^t$	loop
$\bar{b} \rightarrow \bar{s}s\bar{d}$	$K_S K_S$	$\phi K_S$	$(V_{tb}^* V_{td})P^t + (V_{ub}^* V_{ud})P^c$	$\lesssim 1$
$\bar{b} \rightarrow \bar{u}u\bar{d}$	$\pi^+ \pi^-$	$\rho^0 K_S$	$(V_{ub}^* V_{ud})T + (V_{tb}^* V_{td})P^t$	loop
$\bar{b} \rightarrow \bar{c}u\bar{d}$	$D_{CP}\pi^0$	$D_{CP}K_S$	$(V_{cb}^* V_{ud})T + (V_{ub}^* V_{cd})T'$	$\lambda^2$
$\bar{b} \rightarrow \bar{c}u\bar{s}$	$D_{CP}K_S$	$D_{CP}\phi$	$(V_{cb}^* V_{us})T + (V_{ub}^* V_{cs})T'$	$\lesssim 1$

The cleanliness of the theoretical interpretation of  $S_f$  can be assessed from the information in the last column of Table 13.1. In case of small uncertainties, the expression for  $S_f$  in terms of CKM phases can be deduced from the fourth column of Table 13.1 in combination with Eq. (13.90) (and, for  $b \rightarrow q\bar{q}s$  decays, the example in Eq. (13.96)). Here we consider several interesting examples.

For  $B^0 \rightarrow J/\psi K_S$  and other  $\bar{b} \rightarrow \bar{c}c\bar{s}$  processes, we can neglect the  $P^u$  contribution to  $A_f$ , in the Standard Model, to an approximation that is better than one percent, giving:

$$\lambda_{\psi K_S} = -e^{-2i\beta} \Rightarrow S_{\psi K_S} = \sin 2\beta, \quad C_{\psi K_S} = 0. \quad (13.98)$$

It is important to verify experimentally the level of suppression of the penguin contribution. Methods based on flavor symmetries [78–81] allow limits to be obtained. All are currently consistent with the  $P^u$  term being negligible.

In the presence of new physics,  $A_f$  is still likely to be dominated by the  $T$  term, but the mixing amplitude might be modified. We learn that, model-independently,  $C_f \approx 0$  while  $S_f$  cleanly determines the mixing phase ( $\phi_M - 2\arg(V_{cb}^* V_{cd}^*)$ ). The experimental measurement [28],  $S_{\psi K} = +0.691 \pm 0.017$ , gave the first precision test of the Kobayashi-Maskawa mechanism, and its consistency with the predictions for  $\sin 2\beta$  makes it very likely that this mechanism is indeed the dominant source of  $CP$  violation in the quark sector.

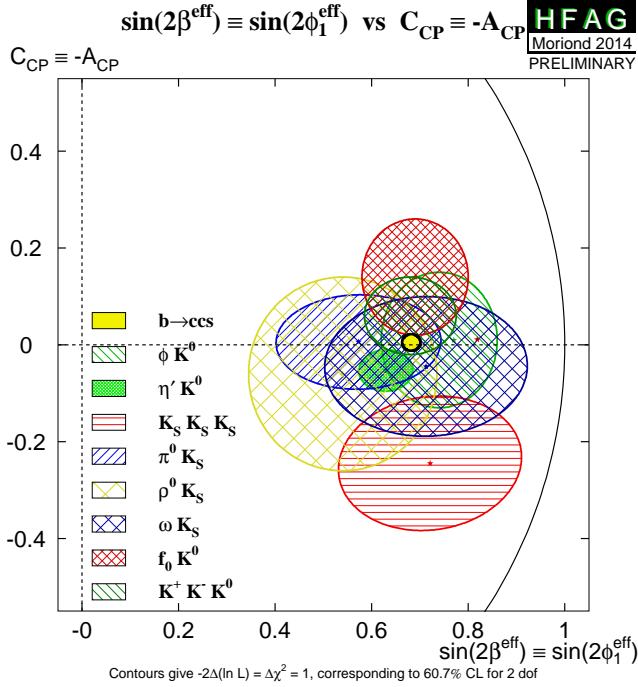
For  $B^0 \rightarrow \phi K_S$  and other  $\bar{b} \rightarrow \bar{s}s\bar{s}$  processes (as well as some  $\bar{b} \rightarrow \bar{u}u\bar{s}$  processes), we can neglect the subdominant contributions, in the Standard Model, to an approximation that is good to the order of a few percent:

$$\lambda_{\phi K_S} = -e^{-2i\beta} \Rightarrow S_{\phi K_S} = \sin 2\beta, \quad C_{\phi K_S} = 0. \quad (13.99)$$

A review of explicit calculations of the effects of subleading amplitudes can be found in Ref. [82]. In the presence of new physics, both  $A_f$  and  $\mathbf{M}_{12}$  can have contributions that are comparable in size to those of the Standard Model and carry new weak phases. Such a situation gives several interesting consequences for penguin-dominated  $b \rightarrow q\bar{q}s$  decays ( $q = u, d, s$ ) to a final state  $f$ :

1. The value of  $-\eta_f S_f$  may be different from  $S_{\psi K_S}$  by more than a few percent, where  $\eta_f$  is the  $CP$  eigenvalue of the final state.
2. The values of  $\eta_f S_f$  for different final states  $f$  may be different from each other by more than a few percent (for example,  $S_{\phi K_S} \neq S_{\eta' K_S}$ ).
3. The value of  $C_f$  may be different from zero by more than a few percent.

While a clear interpretation of such signals in terms of Lagrangian parameters will be difficult because, under these circumstances, hadronic parameters play a role, any of the above three options will clearly signal new physics. Fig. 13.3 summarizes the present experimental results: none of the possible signatures listed above is unambiguously established, but there is definitely still room for new physics.



**Figure 13.3:** Summary of the results [28] of time-dependent analyses of  $b \rightarrow q\bar{q}s$  decays, which are potentially sensitive to new physics.

For the  $\bar{b} \rightarrow \bar{u}u\bar{d}$  process  $B \rightarrow \pi\pi$  and other related channels, the penguin-to-tree ratio can be estimated using SU(3) relations and experimental data on related  $B \rightarrow K\pi$  decays. The result (for  $\pi\pi$ ) is that the suppression is at the level of 0.2–0.3 and so cannot be neglected. The expressions for  $S_{\pi\pi}$  and  $C_{\pi\pi}$  to leading order in  $R_{PT} \equiv (|V_{tb}V_{td}|P_{\pi\pi}^t)/(|V_{ub}V_{ud}|T_{\pi\pi})$  are:

$$\lambda_{\pi\pi} = e^{2i\alpha} \left[ (1 - R_{PT}e^{-i\alpha}) / (1 - R_{PT}e^{+i\alpha}) \right] \Rightarrow$$

$$S_{\pi\pi} \approx \sin 2\alpha + 2 \operatorname{Re}(R_{PT}) \cos 2\alpha \sin \alpha, \quad C_{\pi\pi} \approx 2 \operatorname{Im}(R_{PT}) \sin \alpha. \quad (13.100)$$

Note that  $R_{PT}$  is mode-dependent and, in particular, could be different for  $\pi^+\pi^-$  and  $\pi^0\pi^0$ . If strong phases can be neglected, then  $R_{PT}$  is real, resulting in  $C_{\pi\pi} = 0$ . The size of  $C_{\pi\pi}$  is an indicator of how large the strong phase is. The present experimental average is  $C_{\pi^+\pi^-} = -0.31 \pm 0.05$  [28]. As concerns  $S_{\pi\pi}$ , it is clear from Eq. (13.100) that the relative size or strong phase of the penguin contribution must be known to extract  $\alpha$ . This is the problem of penguin pollution.

The cleanest solution involves isospin relations among the  $B \rightarrow \pi\pi$  amplitudes [83]:

$$\frac{1}{\sqrt{2}} A_{\pi^+\pi^-} + A_{\pi^0\pi^0} = A_{\pi^+\pi^0}. \quad (13.101)$$

The method exploits the fact that the penguin contribution to  $P_{\pi\pi}^t$  is pure  $\Delta I = 1/2$  (this is not true for the electroweak penguins which, however, are expected to be small), while the tree contribution to  $T_{\pi\pi}$  contains pieces that are both  $\Delta I = 1/2$  and  $\Delta I = 3/2$ . A simple geometric construction then allows one to find  $R_{PT}$  and extract  $\alpha$  cleanly from  $S_{\pi^+\pi^-}$ . The key experimental difficulty is that one must measure accurately the separate rates for  $B^0$  and  $\bar{B}^0 \rightarrow \pi^0\pi^0$ .

$CP$  asymmetries in  $B \rightarrow \rho\pi$  and  $B \rightarrow \rho\rho$  can also be used to determine  $\alpha$ . In particular, the  $B \rightarrow \rho\rho$  measurements are presently very significant in constraining  $\alpha$ . The extraction proceeds via isospin analysis similar to that of  $B \rightarrow \pi\pi$ . There are, however, several important differences. First, due to the finite width of the  $\rho$  mesons, a final  $(\rho\rho)_{I=1}$  state is possible [84]. The effect is, however, of the order of  $(\Gamma_\rho/m_\rho)^2 \sim 0.04$ . Second, due to the presence of three helicity states for the two vector mesons, angular analysis is needed to separate the  $CP$ -even and  $CP$ -odd components. The theoretical expectation is that the  $CP$ -odd component is small, which is supported by experiments which find that the  $\rho^+\rho^-$  and  $\rho^\pm\rho^0$  modes are dominantly longitudinally polarized. Third, an important advantage of the  $\rho\rho$  modes is that the penguin contribution is expected to be small due to different hadronic dynamics. This expectation is confirmed by the smallness of  $\mathcal{B}(B^0 \rightarrow \rho^0\rho^0) = (0.96 \pm 0.15) \times 10^{-6}$  [28,85] compared to  $\mathcal{B}(B^0 \rightarrow \rho^+\rho^-) = (24.2 \pm 3.1) \times 10^{-6}$  [28]. Thus,  $S_{\rho^+\rho^-}$  is not far from  $\sin 2\alpha$ . Finally, both  $S_{\rho^0\rho^0}$  and  $C_{\rho^0\rho^0}$  are experimentally accessible, which may allow a precision determination of  $\alpha$ . However, a full isospin analysis should allow that the fractions of longitudinal polarisation in  $B$  and  $\bar{B}$  decays may differ, which has not yet been done by the experiments.

Detailed discussion of the determination of  $\alpha$  with these methods, and the latest world average, can be found in Ref. [34]. The consistency between the range of  $\alpha$  determined by the  $B \rightarrow \pi\pi$ ,  $\rho\pi$  and  $\rho\rho$  measurements and the range allowed by CKM fits (excluding these direct determinations) provides further support to the Kobayashi-Maskawa mechanism.

All modes discussed in this Section so far have possible contributions from penguin amplitudes. As shown in Table 13.1,  $CP$  violation can also be studied with final states, typically containing charmed mesons, where no such contribution is possible. The neutral charmed meson must be reconstructed in a final state, such as a  $CP$  eigenstate, common to  $D^0$  and  $\bar{D}^0$  so that the amplitudes for the  $B$  and  $\bar{B}$  meson decays interfere. Although there is a second tree amplitude with a different weak phase, the contributions of the different diagrams can in many cases be separated experimentally (for example by exploiting different decays of the  $\bar{D}^0$  mesons) making these channels very clean theoretically. The first determination of  $\sin(2\beta)$ , with significance of  $CP$  violation over  $5\sigma$ , with this method has recently been reported [86]. Moreover, the interference between the two tree diagrams gives sensitivity to  $\gamma$ , as will be discussed in Section 13.6.4.

### 13.6.3. $CP$ violation in interference of $B_s^0$ decays with and without mixing :

As discussed in Section 13.6.1, the world average for  $|q/p|$  in the  $B_s^0$  system currently deviates from the Standard Model expectation due to an anomalous value of the dimuon asymmetry. Attributing the dimuon asymmetry result to a fluctuation, we again neglect the deviation of  $|q/p|$  from 1, and use

$$\lambda_f = e^{-i\phi_M(B_s^0)} (\bar{A}_f/A_f). \quad (13.102)$$

Within the Standard Model,

$$e^{-i\phi_M(B_s^0)} = (V_{tb}^* V_{ts}) / (V_{tb} V_{ts}^*). \quad (13.103)$$

Note that  $\Delta\Gamma/\Gamma = 0.122 \pm 0.009$  [28] and therefore  $y$  should not be put to zero in Eqs. (13.35, 13.36). However,  $|q/p| = 1$  is expected to hold to an even better approximation than for  $B^0$  mesons. One therefore obtains

$$\mathcal{A}_f(t) = \frac{S_f \sin(\Delta m t) - C_f \cos(\Delta m t)}{\cosh(\Delta\Gamma t/2) - A_f^{\Delta\Gamma} \sinh(\Delta\Gamma t/2)}, \quad (13.104)$$

$$A_f^{\Delta\Gamma} \equiv \frac{-2 \operatorname{Re}(\lambda_f)}{1 + |\lambda_f|^2}.$$

The presence of the  $A_f^{\Delta\Gamma}$  term implies that information on  $\lambda_f$  can be obtained from analyses that do not use tagging of the initial flavor, through so-called effective lifetime measurements [87].

The  $B_s^0 \rightarrow J/\psi\phi$  decay proceeds via the  $\bar{b} \rightarrow \bar{c}c\bar{s}$  transition. The  $CP$  asymmetry in this mode thus determines (with angular analysis to

disentangle the  $CP$ -even and  $CP$ -odd components of the final state)  $\sin 2\beta_s$ , where  $\beta_s$  is defined in Eq. (13.57) [88]. The  $B_s^0 \rightarrow J/\psi \pi^+ \pi^-$  decay, which has a large contributions from  $J/\psi f_0(980)$  and is assumed to also proceed dominantly via the  $\bar{b} \rightarrow \bar{c} c \bar{s}$  transition, has also been used to determine  $\beta_s$ . In this case no angular analysis is necessary, since the final state has been shown to be dominated by the  $CP$ -even component [89]. A first determination of  $\beta_s$  with  $B_s^0 \rightarrow D_s^+ D_s^-$  decays has also been made [90]. The combination of measurements yields [28]

$$-2\beta_s = 0.015 \pm 0.035, \quad (13.105)$$

consistent with the Standard Model prediction,  $\beta_s = 0.0188 \pm 0.0004$  [18].

The experimental investigation of  $CP$  violation in the  $B_s^0$  sector is still at a relatively early stage, and far fewer modes have been studied than in the  $B^0$  system. First results on the  $\bar{b} \rightarrow \bar{q} q \bar{s}$  decays  $B_s^0 \rightarrow \phi \phi$  and  $K^+ K^-$  have been reported. More channels are expected to be studied in the near future.

#### 13.6.4. Direct $CP$ violation in the $B$ system :

An interesting class of decay modes is that of the tree-level decays  $B^\pm \rightarrow D^{(*)} K^\pm$ . These decays provide golden methods for a clean determination of the angle  $\gamma$  [91–95]. The method uses the decays  $B^+ \rightarrow D^0 K^+$ , which proceeds via the quark transition  $\bar{b} \rightarrow \bar{u} c \bar{s}$ , and  $B^+ \rightarrow \bar{D}^0 K^+$ , which proceeds via the quark transition  $\bar{b} \rightarrow \bar{c} u \bar{s}$ , with the  $D^0$  and  $\bar{D}^0$  decaying into a common final state. The decays into common final states, such  $(\pi^0 K_S) D K^+$ , involve interference effects between the two amplitudes, with sensitivity to the relative phase,  $\delta + \gamma$  ( $\delta$  is the relevant strong phase). The  $CP$ -conjugate processes are sensitive to  $\delta - \gamma$ . Measurements of branching ratios and  $CP$  asymmetries allow the determination of  $\gamma$  and  $\delta$  from amplitude triangle relations. The method suffers from discrete ambiguities but, since all hadronic parameters can be determined from the data, has negligible theoretical uncertainty [96].

Unfortunately, the smallness of the CKM-suppressed  $b \rightarrow u$  transitions makes it difficult at present to use the simplest methods [91–93] to determine  $\gamma$ . These difficulties are overcome (and the discrete ambiguities are removed) by performing a Dalitz plot analysis for multi-body  $D$  decays [94,95].

Constraints on  $\gamma$  from combinations of results on various  $B \rightarrow D^{(*)} K^{(*)}$  processes have been obtained by experiments [97–99], with world averages in Refs. [18,19]. Detailed discussion of the determination of  $\gamma$  with these methods, and the latest world average, can be found in Ref. [34]. The consistency between the range of  $\gamma$  determined by the  $B \rightarrow DK$  measurements and the range allowed by CKM fits (excluding these direct determinations) provides further support to the Kobayashi-Maskawa mechanism. As more data becomes available, determinations of  $\gamma$  from  $B_s^0 \rightarrow D_s^\mp K^\pm$  [100,101] and  $B^0 \rightarrow DK^{*0}$  [102–105] are expected to also give competitive measurements.

Decays to the final state  $K^\mp \pi^\pm$  provided the first observations of direct  $CP$  violation in both  $B^0$  and  $B_s^0$  systems. The asymmetry arises due to interference between tree and penguin diagrams [106], similar to the effect discussed in Section 13.6.2. In principle, measurements of  $\mathcal{A}_{\bar{B}^0 \rightarrow K^- \pi^+}$  and  $\mathcal{A}_{B_s^0 \rightarrow K^+ \pi^-}$  could be used to determine the weak phase difference  $\gamma$ , but lack of knowledge of the relative magnitude and strong phase of the contributing amplitudes limits the achievable precision. The uncertainties on these hadronic parameters can be reduced by exploiting flavor symmetries, which predict a number of relations between asymmetries in different modes. One such relation is that the partial rate differences for  $B^0$  and  $B_s^0$  decays to  $K^\mp \pi^\pm$  are expected to be approximately equal and opposite [107], which is consistent with current data. It is also expected that the partial rate asymmetries for  $\bar{B}^0 \rightarrow K^- \pi^+$  and  $B^- \rightarrow K^- \pi^0$  should be approximately equal; however, the experimental results currently show a significant discrepancy [28]:

$$\mathcal{A}_{\bar{B}^0 \rightarrow K^- \pi^+} = -0.082 \pm 0.006, \quad \mathcal{A}_{B^- \rightarrow K^- \pi^0} = 0.040 \pm 0.021.$$

It is therefore of great interest to understand whether this originates from Standard Model QCD corrections, or whether it is a signature of

new dynamics. Improved tests of a more precise relation between the partial rate differences of all four  $K\pi$  final states [108–111], currently limited by knowledge of the  $CP$  asymmetry in  $\bar{B}^0 \rightarrow K_S \pi^0$  decays, may help to resolve the situation.

It is also of interest to investigate whether similar patterns appear among the  $CP$  violating asymmetries in  $B$  meson decays to final states containing one pseudoscalar and one vector meson. Since the vector resonance decays to two particles, such channels can be studied through Dalitz plot analysis of the three-body final state. Model-independent analyses of  $B^+ \rightarrow K^+ K^- K^+$ ,  $\pi^+ \pi^- K^+$ ,  $\pi^+ \pi^- \pi^+$  and  $K^+ K^- \pi^+$  decays have revealed large  $CP$  violation effects in certain regions of phase space [112]. It remains to be seen whether these are associated to particular resonances or to interference effects, which will be necessary to understand the underlying dynamics.

### 13.7. Summary and Outlook

$CP$  violation has been experimentally established in  $K$  and  $B$  meson decays. A full list of  $CP$  asymmetries that have been measured at a level higher than  $5\sigma$  is given in the introduction to this review. In Section 13.1.4 we introduced three types of  $CP$ -violating effects. Examples of these three types include the following:

1. All three types of  $CP$  violation have been observed in  $K \rightarrow \pi\pi$  decays:

$$\mathcal{R}e(\epsilon') = \frac{1}{6} \left( \left| \frac{\bar{A}_{\pi^0 \pi^0}}{A_{\pi^0 \pi^0}} \right| - \left| \frac{\bar{A}_{\pi^+ \pi^-}}{A_{\pi^+ \pi^-}} \right| \right) = (2.5 \pm 0.4) \times 10^{-6} \quad (\text{I})$$

$$\mathcal{R}e(\epsilon) = \frac{1}{2} \left( 1 - \left| \frac{q}{p} \right| \right) = (1.66 \pm 0.02) \times 10^{-3} \quad (\text{II})$$

$$\mathcal{I}m(\epsilon) = -\frac{1}{2} \mathcal{I}m(\lambda_{(\pi\pi)_{I=0}}) = (1.57 \pm 0.02) \times 10^{-3}. \quad (\text{III})$$

(13.106)

2.  $CP$  violation in decay has been observed in, for example,  $B^0 \rightarrow K^+ \pi^-$  transitions, while  $CP$  violation in interference of decays with and without mixing has been observed in, for example, the  $B^0 \rightarrow J/\psi K_S$  channel:

$$\mathcal{A}_{K^+ \pi^-} = \frac{|\bar{A}_{K^+ \pi^-}/A_{K^+ \pi^-}|^2 - 1}{|\bar{A}_{K^+ \pi^-}/A_{K^+ \pi^-}|^2 + 1} = -0.082 \pm 0.006 \quad (\text{I})$$

$$S_{\psi K} = \mathcal{I}m(\lambda_{\psi K}) = +0.691 \pm 0.017. \quad (\text{III})$$

(13.107)

Based on Standard Model predictions, further observations of  $CP$  violation in  $B^0$ ,  $B^+$  and  $B_s^0$  decays seem likely in the near future, at both LHCb and its upgrade [113,114] as well as the Belle II experiment [115]. The first observation of  $CP$  violation in  $b$  baryons is also likely to be within reach of LHCb. The same experiments have great potential to improve the sensitivity to  $CP$  violation effects in the charm sector, though uncertainty in the Standard Model predictions makes it difficult to forecast whether or not discoveries will be forthcoming. A number of upcoming experiments have potential to make significant progress on rare kaon decays. Observables that are subject to clean theoretical interpretation, such as  $\beta$  from  $S_{\psi K_S}$ ,  $\beta_s$  from  $B_s^0 \rightarrow J/\psi \phi$ ,  $\mathcal{B}(K_L \rightarrow \pi^0 \nu \bar{\nu})$  and  $\gamma$  from  $CP$  violation in  $B \rightarrow DK$  decays, are of particular value for constraining the values of the CKM parameters and probing the flavor sector of extensions to the Standard Model. Progress in lattice QCD calculations is also needed to complement the anticipated experimental results. Other probes of  $CP$  violation now being pursued experimentally include the electric dipole moments of the neutron and electron, and the decays of tau leptons. Additional processes that are likely to play an important role in future  $CP$  studies include top-quark production and decay, Higgs boson decays and neutrino oscillations.

All measurements of  $CP$  violation to date are consistent with the predictions of the Kobayashi-Maskawa mechanism of the Standard Model. In fact, it is now established that the KM mechanism plays a major role in the  $CP$  violation measured in the quark sector.

However, a dynamically-generated matter-antimatter asymmetry of the universe requires additional sources of *CP* violation, and such sources are naturally generated by extensions to the Standard Model. New sources might eventually reveal themselves as small deviations from the predictions of the KM mechanism, or else might not be observable in the quark sector at all, but observable with future probes such as neutrino oscillations or electric dipole moments. The fundamental nature of *CP* violation demands a vigorous search.

A number of excellent reviews of *CP* violation are available [116–122], where the interested reader may find a detailed discussion of the various topics that are briefly reviewed here.

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## References:

1. J.H. Christenson *et al.*, Phys. Rev. Lett. **13**, 138 (1964).
2. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. Lett. **87**, 091801 (2001).
3. K. Abe *et al.* [Belle Collab.], Phys. Rev. Lett. **87**, 091802 (2001).
4. H. Burkhardt *et al.* [NA31 Collab.], Phys. Lett. **B206**, 169 (1988).
5. V. Fanti *et al.* [NA48 Collab.], Phys. Lett. **B465**, 335 (1999).
6. A. Alavi-Harati *et al.* [KTeV Collab.], Phys. Rev. Lett. **83**, 22 (1999).
7. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. Lett. **93**, 131801 (2004).
8. Y. Chao *et al.* [Belle Collab.], Phys. Rev. Lett. **93**, 191802 (2004).
9. A. Poluektov *et al.* [Belle Collab.], Phys. Rev. **D81**, 112002 (2010).
10. P. del Amo Sanchez *et al.* [BABAR Collab.], Phys. Rev. **D82**, 072004 (2010).
11. R. Aaij *et al.* [LHCb Collab.], Phys. Lett. **B712**, 203 (2012).
12. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **110**, 221601 (2013).
13. See results on the “Time reversal invariance,” within the review on “Tests of Conservation Laws,” in this *Review*.
14. J. Bernabeu, F. Martinez-Vidal, and P. Villanueva-Perez, JHEP **08**, 064 (2012).
15. J.P. Lees *et al.* [BABAR Collab.], Phys. Rev. Lett. **109**, 211801 (2012).
16. See, for example, R. F. Streater and A. S. Wightman, *CPT, Spin and Statistics, and All That*, reprinted by Addison-Wesley, New York (1989).
17. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
18. J. Charles *et al.* [CKMfitter Group], Eur. Phys. J. **C41**, 1 (2005); updated results and plots available at: <http://ckmfitter.in2p3.fr>.
19. M. Bona *et al.* [UTfit Collab.], JHEP **10**, 081 (2006); updated results and plots available at: <http://www.utfit.org/UTfit>.
20. A.D. Sakharov, Pisma Zh. Eksp. Teor. Fiz. **5**, 32 (1967) [Sov. Phys. JETP Lett. **5**, 24 (1967)].
21. For a review, see *e.g.*, A. Riotto, “Theories of baryogenesis,” hep-ph/9807454.
22. M. Fukugita and T. Yanagida, Phys. Lett. **B174**, 45 (1986).
23. For a review, see *e.g.*, S. Davidson, E. Nardi and Y. Nir, Phys. Reports **466**, 105 (2008).
24. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. **B716**, 1 (2012).
25. S. Chatrchyan *et al.* [CMS Collab.], Phys. Lett. **B716**, 30 (2012).
26. See the *K*-meson Listings in this *Review*.
27. See the *B*-meson Listings in this *Review*.
28. Y. Amhis *et al.* [HFAG Collab.], arXiv:1412.7515 [hep-ex], and online update at <http://www.slac.stanford.edu/xorg/hfag>.
29. V. Weisskopf and E. P. Wigner, Z. Phys. **63**, 54 (1930); Z. Phys. **65**, 18 (1930) [See also Appendix A of P.K. Kabir, *The CP Puzzle: Strange Decays of the Neutral Kaon*, Academic Press (1968)].
30. See the review on “ $D^0 - \bar{D}^0$  Mixing” in this *Review*.
31. O. Long *et al.*, Phys. Rev. **D68**, 034010 (2003).
32. M. Gronau, Y. Grossman, and J.L. Rosner, Phys. Lett. **B508**, 37 (2001).
33. L. Wolfenstein, Phys. Rev. Lett. **13**, 562 (1964).
34. See the review on “Cabibbo-Kobayashi-Maskawa Mixing Matrix,” in this *Review*.
35. L. Wolfenstein, Phys. Rev. Lett. **51**, 1945 (1983).
36. A.J. Buras, M.E. Lautenbacher, and G. Ostermaier, Phys. Rev. **D50**, 3433 (1994).
37. C. Jarlskog, Phys. Rev. Lett. **55**, 1039 (1985).
38. See the review on “*CP* violation in  $K_S \rightarrow 3\pi$ ,” in this *Review*.
39. T. Blum *et al.*, Phys. Rev. **D91**, 074502 (2015).
40. Z. Bai *et al.*, arXiv:1505.07863 [hep-lat].
41. A.J. Buras *et al.*, JHEP **1511**, 202 (2015).
42. Y. Grossman and Y. Nir, Phys. Lett. **B398**, 163 (1997).
43. L.S. Littenberg, Phys. Rev. **D39**, 3322 (1989).
44. A.J. Buras, Phys. Lett. **B333**, 476 (1994).
45. G. Buchalla and A.J. Buras, Nucl. Phys. **B400**, 225 (1993).
46. J. Brod, M. Gorbahn, and E. Stamou, Phys. Rev. **D83**, 034030 (2011).
47. J.K. Ahn *et al.* [E391a Collab.], Phys. Rev. **D81**, 072004 (2010).
48. T. Yamanaka [KOTO Collab.], Prog. Theor. Exp. Phys. **2012**, 02B006 (2012).
49. M. Mirra [NA62 Collab.], Nuovo Cimento **C038**, 13 (2015).
50. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. Lett. **98**, 211802 (2007).
51. M. Staric *et al.* [Belle Collab.], Phys. Rev. Lett. **98**, 211803 (2007).
52. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. Lett. **100**, 121802 (2008).
53. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **110**, 101802 (2013).
54. See the *D*-meson Listings in this *Review*.
55. Y. Grossman, A. L. Kagan, and Y. Nir, Phys. Rev. **D75**, 036008 (2007).
56. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **108**, 111602 (2012).
57. S. Bergmann *et al.*, Phys. Lett. **B486**, 418 (2000).
58. M. Gersabeck *et al.*, J. Phys. **G39**, 045005 (2012).
59. D.M. Asner *et al.* [CLEO Collab.], Phys. Rev. **D78**, 012001 (2008).
60. M. Ablikim *et al.* [BESIII Collab.], Phys. Lett. **B734**, 227 (2014).
61. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **111**, 251801 (2013).
62. M. Ciuchini *et al.*, Phys. Lett. **B655**, 162 (2007).
63. Y. Grossman, Y. Nir, and G. Perez, Phys. Rev. Lett. **103**, 071602 (2009).
64. A.L. Kagan and M.D. Sokoloff, Phys. Rev. **D80**, 076008 (2009).
65. Y. Grossman and Y. Nir, JHEP **04**, 002 (2012).
66. B.R. Ko *et al.* [Belle Collab.], Phys. Rev. Lett. **109**, 021601 (2012); Erratum-ibid. **109**, 119903 (2012).
67. B. Aubert *et al.* [BABAR Collab.], Phys. Rev. **D78**, 051102 (2008).
68. I. Bediaga *et al.*, Phys. Rev. **D80**, 096006 (2009); Phys. Rev. **D86**, 036005 (2012).
69. M. Williams, Phys. Rev. **D84**, 054015 (2011).
70. A. Lenz and U. Nierste, arXiv:1102.4274 [hep-ph].
71. V.M. Abazov *et al.* [D0 Collab.], Phys. Rev. **D82**, 032001 (2010); Phys. Rev. **D84**, 052007 (2011); Phys. Rev. **D89**, 012002 (2014).
72. A.B. Carter and A.I. Sanda, Phys. Rev. Lett. **45**, 952 (1980); Phys. Rev. **D23**, 1567 (1981).
73. I.I. Bigi and A.I. Sanda, Nucl. Phys. **B193**, 85 (1981).
74. I. Dumietz and J.L. Rosner, Phys. Rev. **D34**, 1404 (1986).
75. Ya.I. Azimov, N.G. Uraltsev, and V.A. Khoze, Sov. J. Nucl. Phys. **45**, 878 (1987) [Yad. Fiz. **45**, 1412 (1987)].
76. I.I. Bigi and A.I. Sanda, Nucl. Phys. **B281**, 41 (1987).
77. G. Buchalla, A.J. Buras, and M.E. Lautenbacher, Rev. Mod. Phys. **68**, 1125 (1996).
78. R. Fleischer, Eur. Phys. J. **C10**, 299 (1999).

79. M. Ciuchini, M. Pierini, and L. Silvestrini, Phys. Rev. Lett. **95**, 221804 (2005).
80. S. Faller *et al.*, Phys. Rev. **D79**, 014030 (2009).
81. M. Jung, Phys. Rev. **D86**, 053008 (2012).
82. L. Silvestrini, Ann. Rev. Nucl. and Part. Sci. **57**, 405 (2007).
83. M. Gronau and D. London, Phys. Rev. Lett. **65**, 3381 (1990).
84. A.F. Falk *et al.*, Phys. Rev. **D69**, 011502 (2004).
85. R. Aaij *et al.* [LHCb Collab.], Phys. Lett. **B747**, 468 (2015).
86. A. Abdesselam *et al.* [BABAR and Belle Collabs.], Phys. Rev. Lett. **115**, 121604 (2015).
87. R. Fleischer and R. Kneijens, Eur. Phys. J. **C71**, 1789 (2011).
88. A.S. Dighe, I. Dunietz, and R. Fleischer, Eur. Phys. J. **C6**, 647 (1999).
89. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. **D89**, 092006 (2014).
90. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **113**, 211801 (2014).
91. M. Gronau and D. London, Phys. Lett. **B253**, 483 (1991).
92. M. Gronau and D. Wyler, Phys. Lett. **B265**, 172 (1991).
93. D. Atwood, I. Dunietz, and A. Soni, Phys. Rev. Lett. **78**, 3257 (1997).
94. D. Atwood, I. Dunietz, and A. Soni, Phys. Rev. **D63**, 036005 (2001).
95. A. Giri *et al.*, Phys. Rev. **D68**, 054018 (2003).
96. J. Brod and J. Zupan, JHEP **01**, 051 (2014).
97. J. P. Lees *et al.* [BaBar Collab.], Phys. Rev. **D87**, 052015 (2013).
98. K. Trabelsi [Belle Collab.], [arXiv:1301.2033 \[hep-ex\]](#).
99. R. Aaij *et al.* [LHCb Collab.], Phys. Lett. **B726**, 151 (2013), LHCb-CONF-2014-004.
100. R. Aleksan, I. Dunietz, and B. Kayser, Z. Phys. **C54**, 653 (1992).
101. R. Fleischer, Nucl. Phys. **B671**, 459 (2003).
102. I. Dunietz, Phys. Lett. **B270**, 75 (1991).
103. M. Gronau, Phys. Lett. **B557**, 198 (2003).
104. T. Gershon, Phys. Rev. **D79**, 051301 (2009).
105. T. Gershon and M. Williams, Phys. Rev. **D80**, 092002 (2009).
106. M. Bander, D. Silverman, and A. Soni, Phys. Rev. Lett. **43**, 242 (1979).
107. X.-G. He, Eur. Phys. J. **C9**, 443 (1999).
108. D. Atwood and A. Soni, Phys. Rev. **D58**, 036005 (1998).
109. M. Gronau and J. L. Rosner, Phys. Rev. **D59**, 113002 (1999).
110. H.J. Lipkin, Phys. Lett. **B445**, 403 (1999).
111. M. Gronau, Phys. Lett. **B627**, 82 (2005).
112. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **111**, 101801 (2013); Phys. Rev. Lett. **112**, 011801 (2014); Phys. Rev. **D90**, 112004 (2014).
113. A.A. Alves *et al.* [LHCb Collab.], JINST **3** S08005 (2008).
114. I. Bediaga *et al.* [LHCb Collab.], CERN-LHCC-2012-007, LHCb-TDR-12.
115. T. Aushev *et al.*, [arXiv:1002.5012 \[hep-ex\]](#), KEK Report 2009-12.
116. G.C. Branco, L. Lavoura, and J.P. Silva, *CP Violation*, Oxford University Press, Oxford (1999).
117. I.I.Y. Bigi and A.I. Sanda, *CP Violation*, Cambridge Monogr., Part. Phys. Nucl. Phys. Cosmol. **9**, 1 (2000).
118. A.J. Bevan *et al.* [BABAR and Belle Collabs.], Eur. Phys. J. **C74**, 3026 (2014).
119. H.R. Quinn and Y. Nir, *"The Mystery of the Missing Antimatter,"* Princeton University Press, Princeton (2008).
120. T.E. Browder *et al.*, Rev. Mod. Phys. **81**, 1887 (2009).
121. M. Ciuchini and A. Stocchi, Ann. Rev. Nucl. and Part. Sci. **61**, 491 (2011).
122. R. Aaij *et al.* [LHCb Collab.] and A. Bharucha *et al.*, Eur. Phys. J. **C73**, 2373 (2013).

## 14. NEUTRINO MASS, MIXING, AND OSCILLATIONS

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The experiments with solar, atmospheric, reactor and accelerator neutrinos have provided compelling evidences for oscillations of neutrinos caused by nonzero neutrino masses and neutrino mixing. The data imply the existence of 3-neutrino mixing in vacuum. We review the theory of neutrino oscillations, the phenomenology of neutrino mixing, the problem of the nature - Dirac or Majorana, of massive neutrinos, the issue of CP violation in the lepton sector, and the current data on the neutrino masses and mixing parameters. The open questions and the main goals of future research in the field of neutrino mixing and oscillations are outlined.

### 14.1. Introduction: Massive neutrinos and neutrino mixing

It is a well-established experimental fact that the neutrinos and antineutrinos which take part in the standard charged current (CC) and neutral current (NC) weak interaction are of three varieties (types) or flavours: electron,  $\nu_e$  and  $\bar{\nu}_e$ , muon,  $\nu_\mu$  and  $\bar{\nu}_\mu$ , and tauon,  $\nu_\tau$  and  $\bar{\nu}_\tau$ . The notion of neutrino type or flavour is dynamical:  $\nu_e$  is the neutrino which is produced with  $e^+$ , or produces an  $e^-$ , in CC weak interaction processes;  $\nu_\mu$  is the neutrino which is produced with  $\mu^+$ , or produces  $\mu^-$ , etc. The flavour of a given neutrino is Lorentz invariant. Among the three different flavour neutrinos and antineutrinos, no two are identical. Correspondingly, the states which describe different flavour neutrinos must be orthogonal (within the precision of the current data):  $\langle \nu_l | \nu_l \rangle = \delta_{ll}$ ,  $\langle \bar{\nu}_l | \bar{\nu}_l \rangle = \delta_{ll}$ ,  $\langle \bar{\nu}_l | \nu_l \rangle = 0$ .

It is also well-known from the existing data (all neutrino experiments were done so far with relativistic neutrinos or antineutrinos), that the flavour neutrinos  $\nu_l$  (antineutrinos  $\bar{\nu}_l$ ), are always produced in weak interaction processes in a state that is predominantly left-handed (LH) (right-handed (RH)). To account for this fact,  $\nu_l$  and  $\bar{\nu}_l$  are described in the Standard Model (SM) by a chiral LH flavour neutrino field  $\nu_{lL}(x)$ ,  $l = e, \mu, \tau$ . For massless  $\nu_l$ , the state of  $\nu_l$  ( $\bar{\nu}_l$ ) which the field  $\nu_{lL}(x)$  annihilates (creates) is with helicity  $(-1/2)$  (helicity  $+1/2$ ). If  $\nu_l$  has a non-zero mass  $m(\nu_l)$ , the state of  $\nu_l$  ( $\bar{\nu}_l$ ) is a linear superposition of the helicity  $(-1/2)$  and  $(+1/2)$  states, but the helicity  $+1/2$  state (helicity  $(-1/2)$  state) enters into the superposition with a coefficient  $\propto m(\nu_l)/E$ ,  $E$  being the neutrino energy, and thus is strongly suppressed. Together with the LH charged lepton field  $l_L(x)$ ,  $\nu_{lL}(x)$  forms an  $SU(2)_L$  doublet. In the absence of neutrino mixing and zero neutrino masses,  $\nu_{lL}(x)$  and  $l_L(x)$  can be assigned one unit of the additive lepton charge  $L_l$  and the three charges  $L_l$ ,  $l = e, \mu, \tau$ , are conserved by the weak interaction.

At present there is no compelling evidence for the existence of states of relativistic neutrinos (antineutrinos), which are predominantly right-handed,  $\nu_R$  (left-handed,  $\bar{\nu}_L$ ). If RH neutrinos and LH antineutrinos exist, their interaction with matter should be much weaker than the weak interaction of the flavour LH neutrinos  $\nu_l$  and RH antineutrinos  $\bar{\nu}_l$ , i.e.,  $\nu_R$  ( $\bar{\nu}_L$ ) should be “sterile” or “inert” neutrinos (antineutrinos) [1]. In the formalism of the Standard Model, the sterile  $\nu_R$  and  $\bar{\nu}_L$  can be described by  $SU(2)_L$  singlet RH neutrino fields  $\nu_R(x)$ . In this case,  $\nu_R$  and  $\bar{\nu}_L$  will have no gauge interactions, i.e., will not couple to the weak  $W^\pm$  and  $Z^0$  bosons. If present in an extension of the Standard Model, the RH neutrinos can play a crucial role i) in the generation of neutrino masses and mixing, ii) in understanding the remarkable disparity between the magnitudes of neutrino masses and the masses of the charged leptons and quarks, and iii) in the generation of the observed matter-antimatter asymmetry of the Universe (via the leptogenesis mechanism [2]). In this scenario which is based on the see-saw theory [3], there is a link between the generation of neutrino masses and the generation of the baryon asymmetry of the Universe. The simplest hypothesis (based on symmetry considerations) is that to each LH flavour neutrino field  $\nu_{lL}(x)$  there corresponds a RH neutrino field  $\nu_{lR}(x)$ ,  $l = e, \mu, \tau$ , although schemes with less (more) than three RH neutrinos are also being considered.

The experiments with solar, atmospheric, reactor and accelerator neutrinos have provided compelling evidences for the existence of neutrino oscillations [4,5], transitions in flight between the different flavour neutrinos  $\nu_e$ ,  $\nu_\mu$ ,  $\nu_\tau$  (antineutrinos  $\bar{\nu}_e$ ,  $\bar{\nu}_\mu$ ,  $\bar{\nu}_\tau$ ), caused by nonzero neutrino masses and neutrino mixing. The existence of flavour neutrino oscillations implies that if a neutrino of a given flavour, say  $\nu_\mu$ , with energy  $E$  is produced in some weak interaction process, at a sufficiently large distance  $L$  from the  $\nu_\mu$  source the probability to find a neutrino of a different flavour, say  $\nu_\tau$ ,  $P(\nu_\mu \rightarrow \nu_\tau; E, L)$ , is different from zero.  $P(\nu_\mu \rightarrow \nu_\tau; E, L)$  is called the  $\nu_\mu \rightarrow \nu_\tau$  oscillation or transition probability. If  $P(\nu_\mu \rightarrow \nu_\tau; E, L) \neq 0$ , the probability that  $\nu_\mu$  will not change into a neutrino of a different flavour, i.e., the “ $\nu_\mu$  survival probability”  $P(\nu_\mu \rightarrow \nu_\mu; E, L)$ , will be smaller than one. If only muon neutrinos  $\nu_\mu$  are detected in a given experiment and they take part in oscillations, one would observe a “disappearance” of muon neutrinos on the way from the  $\nu_\mu$  source to the detector. Disappearance of the solar  $\nu_e$ , reactor  $\bar{\nu}_e$  and of atmospheric  $\nu_\mu$  and  $\bar{\nu}_\mu$  due to the oscillations have been observed respectively, in the solar neutrino [6–14], KamLAND [15,16] and Super-Kamiokande [17,18] experiments. Strong evidences for  $\nu_\mu$  disappearance due to oscillations were obtained also in the long-baseline accelerator neutrino experiments K2K [19]. Subsequently, the MINOS [20,21] and T2K [22,23] long baseline experiments reported compelling evidence for  $\nu_\mu$  disappearance due to oscillations, while evidences for  $\nu_\tau$  appearance due to  $\nu_\mu \rightarrow \nu_\tau$  oscillations were published by the Super-Kamiokande [24] and OPERA [25] collaborations. As a consequence of the results of the experiments quoted above the existence of oscillations or transitions of the solar  $\nu_e$ , atmospheric  $\nu_\mu$  and  $\bar{\nu}_\mu$ , accelerator  $\nu_\mu$  (at  $L \sim 250$  km,  $L \sim 295$  km and  $L \sim 730$  km) and reactor  $\bar{\nu}_e$  (at  $L \sim 180$  km), driven by nonzero neutrino masses and neutrino mixing, was firmly established. There are strong indications that the solar  $\nu_e$  transitions are affected by the solar matter [26,27].

Further important developments took place in the period starting from June 2011. First, the T2K Collaboration reported [28] indications for  $\nu_\mu \rightarrow \nu_e$  oscillations, i.e., of “appearance” of  $\nu_e$  in a beam of  $\nu_\mu$ , which had a statistical significance of  $2.5\sigma$ . The MINOS [29] Collaboration also obtained data consistent with  $\nu_\mu \rightarrow \nu_e$  oscillations. Subsequently, the Double Chooz Collaboration reported [30] indications for disappearance of reactor  $\bar{\nu}_e$  at  $L \sim 1.1$  km. Strong evidences for reactor  $\bar{\nu}_e$  disappearance at  $L \sim 1.65$  km and  $L \sim 1.38$  km and (with statistical significance of  $5.2\sigma$  and  $4.9\sigma$ ) were obtained respectively in the Daya Bay [31] and RENO [32] experiments. Further evidences for reactor  $\bar{\nu}_e$  disappearance (at  $2.9\sigma$ ) and for  $\nu_\mu \rightarrow \nu_e$  oscillations (at  $3.1\sigma$ ) were reported by the Double Chooz [33] and T2K [34] experiments, while the Daya Bay and RENO Collaborations presented updated, more precise results on reactor  $\bar{\nu}_e$  disappearance [35,36,37] (for the latest results of the Daya Bay [38], RENO [39], Double Chooz [40], MINOS [41] and T2K experiments [42], see Section 14.12).

Oscillations of neutrinos are a consequence of the presence of flavour neutrino mixing, or lepton mixing, in vacuum. In the formalism of local quantum field theory, used to construct the Standard Model, this means that the LH flavour neutrino fields  $\nu_{lL}(x)$ , which enter into the expression for the lepton current in the CC weak interaction Lagrangian, are linear combinations of the fields of three (or more) neutrinos  $\nu_j$ , having masses  $m_j \neq 0$ :

$$\nu_{lL}(x) = \sum_j U_{lj} \nu_{jL}(x), \quad l = e, \mu, \tau, \quad (14.1)$$

where  $\nu_{jL}(x)$  is the LH component of the field of  $\nu_j$  possessing a mass  $m_j$  and  $U$  is a unitary matrix - the neutrino mixing matrix [1,4,5]. The matrix  $U$  is often called the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) or Maki-Nakagawa-Sakata (MNS) mixing matrix. Obviously, Eq. (14.1) implies that the individual lepton charges  $L_l$ ,  $l = e, \mu, \tau$ , are not conserved.

All compelling neutrino oscillation data can be described assuming 3-flavour neutrino mixing in vacuum. The data on the invisible decay width of the  $Z$ -boson is compatible with only 3 light flavour neutrinos coupled to  $Z$  [43]. The number of massive neutrinos  $\nu_j$ ,

$n$ , can, in general, be bigger than 3,  $n > 3$ , if, for instance, there exist sterile neutrinos and they mix with the flavour neutrinos. It is firmly established on the basis of the current data that at least 3 of the neutrinos  $\nu_j$ , say  $\nu_1, \nu_2, \nu_3$ , must be light,  $m_{1,2,3} \lesssim 1$  eV (Section 14.12), and must have different masses,  $m_1 \neq m_2 \neq m_3$ . At present there are several experimental hints for existence of one or two light sterile neutrinos at the eV scale, which mix with the flavour neutrinos, implying the presence in the neutrino mixing of additional one or two neutrinos,  $\nu_4$  or  $\nu_{4,5}$ , with masses  $m_4$  ( $m_{4,5}$ )  $\sim 1$  eV. These hints will be briefly discussed in Section 14.13 of the present review.

Being electrically neutral, the neutrinos with definite mass  $\nu_j$  can be Dirac fermions or Majorana particles [44,45]. The first possibility is realized when there exists a lepton charge carried by the neutrinos  $\nu_j$ , which is conserved by the particle interactions. This could be, *e.g.*, the total lepton charge  $L = L_e + L_\mu + L_\tau$ :  $L(\nu_j) = 1$ ,  $j = 1, 2, 3$ . In this case the neutrino  $\nu_j$  has a distinctive antiparticle  $\bar{\nu}_j$ :  $\bar{\nu}_j$  differs from  $\nu_j$  by the value of the lepton charge  $L$  it carries,  $L(\bar{\nu}_j) = -1$ . The massive neutrinos  $\nu_j$  can be Majorana particles if no lepton charge is conserved (see, *e.g.*, Refs. [46,47]). A massive Majorana particle  $\chi_j$  is identical with its antiparticle  $\bar{\chi}_j$ :  $\chi_j \equiv \bar{\chi}_j$ . On the basis of the existing neutrino data it is impossible to determine whether the massive neutrinos are Dirac or Majorana fermions.

In the case of  $n$  neutrino flavours and  $n$  massive neutrinos, the  $n \times n$  unitary neutrino mixing matrix  $U$  can be parametrized by  $n(n-1)/2$  Euler angles and  $n(n+1)/2$  phases. If the massive neutrinos  $\nu_j$  are Dirac particles, only  $(n-1)(n-2)/2$  phases are physical and can be responsible for CP violation in the lepton sector. In this respect the neutrino (lepton) mixing with Dirac massive neutrinos is similar to the quark mixing. For  $n = 3$  there is just one CP violating phase in  $U$ , which is usually called “the Dirac CP violating phase.” CP invariance holds if (in a certain standard convention)  $U$  is real,  $U^* = U$ .

If, however, the massive neutrinos are Majorana fermions,  $\nu_j \equiv \chi_j$ , the neutrino mixing matrix  $U$  contains  $n(n-1)/2$  CP violation phases [48,49], *i.e.*, by  $(n-1)$  phases more than in the Dirac neutrino case: in contrast to Dirac fields, the massive Majorana neutrino fields cannot “absorb” phases. In this case  $U$  can be cast in the form [48]

$$U = V P \quad (14.2)$$

where the matrix  $V$  contains the  $(n-1)(n-2)/2$  Dirac CP violation phases, while  $P$  is a diagonal matrix with the additional  $(n-1)$  Majorana CP violation phases  $\alpha_{21}, \alpha_{31}, \dots, \alpha_{n1}$ ,

$$P = \text{diag} \left( 1, e^{i\frac{\alpha_{21}}{2}}, e^{i\frac{\alpha_{31}}{2}}, \dots, e^{i\frac{\alpha_{n1}}{2}} \right). \quad (14.3)$$

The Majorana phases will conserve CP if [50]  $\alpha_{j1} = \pi q_j$ ,  $q_j = 0, 1, 2$ ,  $j = 2, 3, \dots, n$ . In this case  $\exp[i(\alpha_{j1} - \alpha_{k1})] = \pm 1$  has a simple physical interpretation: this is the relative CP-parity of Majorana neutrinos  $\chi_j$  and  $\chi_k$ . The condition of CP invariance of the leptonic CC weak interaction in the case of mixing and massive Majorana neutrinos reads [46]:

$$U_{lj}^* = U_{lj} \rho_j, \quad \rho_j = \frac{1}{i} \eta_{CP}(\chi_j) = \pm 1, \quad (14.4)$$

where  $\eta_{CP}(\chi_j) = i\rho_j = \pm i$  is the CP parity of the Majorana neutrino  $\chi_j$  [50]. Thus, if CP invariance holds, the elements of  $U$  are either real or purely imaginary.

In the case of  $n = 3$  there are altogether 3 CP violation phases - one Dirac and two Majorana. Even in the mixing involving only 2 massive Majorana neutrinos there is one physical CP violation Majorana phase. In contrast, the CC weak interaction is automatically CP-invariant in the case of mixing of two massive Dirac neutrinos or of two quarks.

## 14.2. The three neutrino mixing

All existing compelling data on neutrino oscillations can be described assuming 3-flavour neutrino mixing in vacuum. This is the minimal neutrino mixing scheme which can account for the currently available data on the oscillations of the solar ( $\nu_e$ ), atmospheric ( $\nu_\mu$  and  $\bar{\nu}_\mu$ ), reactor ( $\bar{\nu}_e$ ) and accelerator ( $\nu_\mu$  and  $\bar{\nu}_\mu$ ) neutrinos. The (left-handed) fields of the flavour neutrinos  $\nu_e, \nu_\mu$  and  $\nu_\tau$  in the expression for the weak charged lepton current in the CC weak interaction Lagrangian, are linear combinations of the LH components of the fields of three massive neutrinos  $\nu_j$ :

$$\mathcal{L}_{CC} = -\frac{g}{\sqrt{2}} \sum_{l=e,\mu,\tau} \bar{l}_L(x) \gamma_\alpha \nu_{lL}(x) W^{\alpha\dagger}(x) + h.c.,$$

$$\nu_{lL}(x) = \sum_{j=1}^3 U_{lj} \nu_{jL}(x), \quad (14.5)$$

where  $U$  is the  $3 \times 3$  unitary neutrino mixing matrix [4,5]. As we have discussed in the preceding Section, the mixing matrix  $U$  can be parameterized by 3 angles, and, depending on whether the massive neutrinos  $\nu_j$  are Dirac or Majorana particles, by 1 or 3 CP violation phases [48,49]:

$$U = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \times \text{diag}(1, e^{i\frac{\alpha_{21}}{2}}, e^{i\frac{\alpha_{31}}{2}}). \quad (14.6)$$

where  $c_{ij} = \cos \theta_{ij}$ ,  $s_{ij} = \sin \theta_{ij}$ , the angles  $\theta_{ij} = [0, \pi/2]$ ,  $\delta = [0, 2\pi]$  is the Dirac CP violation phase and  $\alpha_{21}, \alpha_{31}$  are two Majorana CP violation (CPV) phases. Thus, in the case of massive Dirac neutrinos, the neutrino mixing matrix  $U$  is similar, in what concerns the number of mixing angles and CPV phases, to the CKM quark mixing matrix. The presence of two additional physical CPV phases in  $U$  if  $\nu_j$  are Majorana particles is a consequence of the special properties of the latter (see, *e.g.*, Refs. [46,48]).

As we see, the fundamental parameters characterizing the 3-neutrino mixing are: i) the 3 angles  $\theta_{12}, \theta_{23}, \theta_{13}$ , ii) depending on the nature of massive neutrinos  $\nu_j$  - 1 Dirac ( $\delta$ ), or 1 Dirac + 2 Majorana ( $\delta, \alpha_{21}, \alpha_{31}$ ), CPV phases, and iii) the 3 neutrino masses,  $m_1, m_2, m_3$ . Thus, depending on whether the massive neutrinos are Dirac or Majorana particles, this makes 7 or 9 additional parameters in the minimally extended Standard Model of particle interactions with massive neutrinos.

The angles  $\theta_{12}, \theta_{23}$  and  $\theta_{13}$  can be defined via the elements of the neutrino mixing matrix:

$$c_{12}^2 \equiv \cos^2 \theta_{12} = \frac{|U_{e1}|^2}{1 - |U_{e3}|^2}, \quad s_{12}^2 \equiv \sin^2 \theta_{12} = \frac{|U_{e2}|^2}{1 - |U_{e3}|^2}, \quad (14.7)$$

$$s_{13}^2 \equiv \sin^2 \theta_{13} = |U_{e3}|^2, \quad s_{23}^2 \equiv \sin^2 \theta_{23} = \frac{|U_{\mu 3}|^2}{1 - |U_{e3}|^2},$$

$$c_{23}^2 \equiv \cos^2 \theta_{23} = \frac{|U_{\tau 3}|^2}{1 - |U_{e3}|^2}. \quad (14.8)$$

The neutrino oscillation probabilities depend (Section 14.7), in general, on the neutrino energy,  $E$ , the source-detector distance  $L$ , on the elements of  $U$  and, for relativistic neutrinos used in all neutrino experiments performed so far, on  $\Delta m_{ij}^2 \equiv (m_i^2 - m_j^2)$ ,  $i \neq j$ . In the case of 3-neutrino mixing there are only two independent neutrino mass squared differences, say  $\Delta m_{21}^2 \neq 0$  and  $\Delta m_{31}^2 \neq 0$ . The numbering of massive neutrinos  $\nu_j$  is arbitrary. It proves convenient from the point of view of relating the mixing angles  $\theta_{12}, \theta_{23}$  and  $\theta_{13}$  to observables, to identify  $|\Delta m_{21}^2|$  with the smaller of the two neutrino mass squared differences, which, as it follows from the data, is responsible for the solar  $\nu_e$  and, the observed by KamLAND, reactor  $\bar{\nu}_e$  oscillations. We will number (just for convenience) the massive neutrinos in such a way that  $m_1 < m_2$ , so that  $\Delta m_{21}^2 > 0$ . With these choices made, there are two possibilities: either  $m_1 < m_2 < m_3$ ,



or  $m_3 < m_1 < m_2$ . Then the larger neutrino mass square difference  $|\Delta m_{31}^2|$  or  $|\Delta m_{32}^2|$ , can be associated with the experimentally observed oscillations of the atmospheric and accelerator  $\nu_\mu$  and  $\bar{\nu}_\mu$ , as well as of the reactor  $\bar{\nu}_e$  at  $L \sim 1$  km. The effects of  $\Delta m_{31}^2$  or  $\Delta m_{32}^2$  in the oscillations of solar  $\nu_e$ , and of  $\Delta m_{21}^2$  in the oscillations of atmospheric and accelerator  $\nu_\mu$  and  $\bar{\nu}_\mu$  or of the reactor  $\bar{\nu}_e$  at  $L \sim 1$  km, are relatively small and subdominant as a consequence of the facts that i)  $L$ ,  $E$  and  $L/E$  in the experiments with solar  $\nu_e$  and with atmospheric and accelerator  $\nu_\mu$  and  $\bar{\nu}_\mu$ , or with reactor  $\bar{\nu}_e$  and baseline  $L \sim 1$  km, are very different, ii) the conditions of production and propagation (on the way to the detector) of the solar  $\nu_e$  and of the atmospheric or accelerator  $\nu_\mu$  and  $\bar{\nu}_\mu$  and of the reactor  $\bar{\nu}_e$ , are very different, and iii)  $|\Delta m_{21}^2|$  and  $|\Delta m_{31}^2|$  ( $|\Delta m_{32}^2|$ ) in the case of  $m_1 < m_2 < m_3$  ( $m_3 < m_1 < m_2$ ), as it follows from the data, differ by approximately a factor of 30,  $|\Delta m_{21}^2| \ll |\Delta m_{31(32)}^2|$ ,  $|\Delta m_{21}^2|/|\Delta m_{31(32)}^2| \cong 0.03$ . This implies that in both cases of  $m_1 < m_2 < m_3$  and  $m_3 < m_1 < m_2$  we have  $\Delta m_{32}^2 \cong \Delta m_{31}^2$  with  $|\Delta m_{31}^2 - \Delta m_{32}^2| = |\Delta m_{21}^2| \ll |\Delta m_{31,32}^2|$ . Obviously, in the case of  $m_1 < m_2 < m_3$  ( $m_3 < m_1 < m_2$ ) we have  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ).

It followed from the results of the Chooz experiment [51] with reactor  $\bar{\nu}_e$  and from the more recent data of the Daya Bay, RENO, Double Chooz and T2K experiments (which will be discussed in Section 14.12), that, in the convention we use, in which  $0 < \Delta m_{21}^2 < |\Delta m_{31(32)}^2|$ , the element  $|U_{e3}| = \sin \theta_{13}$  of the neutrino mixing matrix  $U$  is relatively small. This makes it possible to identify the angles  $\theta_{12}$  and  $\theta_{23}$  as the neutrino mixing angles associated with the solar  $\nu_e$  and the dominant atmospheric  $\nu_\mu$  (and  $\bar{\nu}_\mu$ ) oscillations, respectively. The angles  $\theta_{12}$  and  $\theta_{23}$  are sometimes called “solar” and “atmospheric” neutrino mixing angles, and are sometimes denoted as  $\theta_{12} = \theta_\odot$  and  $\theta_{23} = \theta_A$  (or  $\theta_{\text{atm}}$ ), while  $\Delta m_{21}^2$  and  $\Delta m_{31}^2$  are often referred to as the “solar” and “atmospheric” neutrino mass squared differences and are often denoted as  $\Delta m_{21}^2 \equiv \Delta m_\odot^2$ ,  $\Delta m_{31}^2 \equiv \Delta m_A^2$  (or  $\Delta m_{\text{atm}}^2$ ).

The solar neutrino data tell us that  $\Delta m_{21}^2 \cos 2\theta_{12} > 0$ . In the convention employed by us we have  $\Delta m_{21}^2 > 0$ . Correspondingly, in this convention one must have  $\cos 2\theta_{12} > 0$ .

Global analyses of the neutrino oscillation data [52,53,54] available by the second half of 2014 allowed us to determine the 3-neutrino oscillation parameters  $\Delta m_{21}^2$ ,  $\theta_{12}$ ,  $|\Delta m_{31}^2|$  ( $|\Delta m_{32}^2|$ ),  $\theta_{23}$  and  $\theta_{13}$  with a relatively high precision.

The authors of the three independent analyses [52,53,54] report practically the same (within  $1\sigma$ ) results on  $\Delta m_{21}^2$ ,  $\sin^2 \theta_{12}$ ,  $|\Delta m_{31}^2|$  and  $\sin^2 \theta_{13}$ . The results obtained in Ref. 52 on  $\sin^2 \theta_{23}$  show, in particular, that for  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ), i.e., for  $m_1 < m_2 < m_3$  ( $m_3 < m_1 < m_2$ ), the best fit value of  $\sin^2 \theta_{23} = 0.437$  (0.455). At the same time, the best fit values of  $\sin^2 \theta_{23}$  reported for  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ) in Ref. 53 and in Ref. 54 read, respectively:  $\sin^2 \theta_{23} = 0.452$  (0.579) and  $\sin^2 \theta_{23} = 0.567$  (0.573). It should be added, however, that the global minima of the corresponding  $\chi^2$  functions in all three cases of the analyses [52,53,54] are accompanied by local minima which are less than  $1\sigma$  away in the value of the  $\chi^2$  function from the corresponding global minima. Taking into account the global and the local minima of the  $\chi^2$  function found in [52,53,54] makes the results on  $\sin^2 \theta_{23}$  (including the  $3\sigma$  allowed ranges) obtained in Refs. [52], [53] and [54] compatible. This, in particular, reflects the fact that the value of  $\sin^2 \theta_{23}$  is still determined experimentally with a relatively large uncertainty.

In all three analyses [52,53,54] the authors find that the best fit value of the Dirac CPV phases  $\delta \cong 3\pi/2$ . According to Ref. 52, the CP conserving values  $\delta = 0$  ( $2\pi$ ) and  $\pi$  ( $\delta = 0$  ( $2\pi$ )) are disfavored at  $1.6\sigma$  to  $2.0\sigma$  (at  $2.0\sigma$ ) for  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ). In the case of  $\Delta m_{31(32)}^2 < 0$ , the value  $\delta = \pi$  is statistically  $1\sigma$  away from the best fit value  $\delta \cong 3\pi/2$ . Similar results are obtained in [53,54].

In August 2015 the first results of the NO $\nu$ A neutrino oscillation experiment were announced [55,56]. These results together with the latest neutrino and the first antineutrino data from the T2K experiment [57,58] (see also Ref. 59) were included, in particular, in the latest analysis of the global neutrino oscillation data performed in Ref. 60. Thus, in Ref. 60 the authors updated the results

obtained earlier in [52,53,54]. We present in Table 14.1 the best fit values and the 99.73% CL allowed ranges of the neutrino oscillation parameters found in Ref. 60 using, in particular, the more “conservative” LID NO $\nu$ A data from Ref. 56. The best fit value of  $\sin^2 \theta_{23}$  found for  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ) in Ref. 60 reads:  $\sin^2 \theta_{23} = 0.437$  (0.569). The authors of Ref. 60 also find that the hint for  $\delta \cong 3\pi/2$  is strengthened by the NO $\nu$ A  $\nu_\mu \rightarrow \nu_e$  and T2K  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation data. The values of  $\delta = \pi/2$  and  $\delta = 0$  ( $2\pi$ ) are disfavored at  $3\sigma$  CL and  $2\sigma$  CL, respectively, while  $\delta = \pi$  is allowed at approximately  $1.6\sigma$  CL ( $1.2\sigma$  CL) for  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ).

**Table 14.1:** The best-fit values and  $3\sigma$  allowed ranges of the 3-neutrino oscillation parameters, derived from a global fit of the current neutrino oscillation data (from [60]). For the Dirac phase  $\delta$  we give the best fit value and the  $2\sigma$  allowed ranges; at  $3\sigma$  no physical values of  $\delta$  are disfavored. The values (values in brackets) correspond to  $m_1 < m_2 < m_3$  ( $m_3 < m_1 < m_2$ ). The definition of  $\Delta m^2$  used is:  $\Delta m^2 = m_3^2 - (m_2^2 + m_1^2)/2$ . Thus,  $\Delta m^2 = \Delta m_{31}^2 - \Delta m_{21}^2/2 > 0$ , if  $m_1 < m_2 < m_3$ , and  $\Delta m^2 = \Delta m_{32}^2 + \Delta m_{21}^2/2 < 0$  for  $m_3 < m_1 < m_2$ .

Parameter	best-fit	$3\sigma$
$\Delta m_{21}^2$ [ $10^{-5}$ eV $^2$ ]	7.37	6.93 – 7.97
$ \Delta m^2 $ [ $10^{-3}$ eV $^2$ ]	2.50 (2.46)	2.37 – 2.63 (2.33 – 2.60)
$\sin^2 \theta_{12}$	0.297	0.250 – 0.354
$\sin^2 \theta_{23}$ , $\Delta m^2 > 0$	0.437	0.379 – 0.616
$\sin^2 \theta_{23}$ , $\Delta m^2 < 0$	0.569	0.383 – 0.637
$\sin^2 \theta_{13}$ , $\Delta m^2 > 0$	0.0214	0.0185 – 0.0246
$\sin^2 \theta_{13}$ , $\Delta m^2 < 0$	0.0218	0.0186 – 0.0248
$\delta/\pi$	1.35 (1.32)	(0.92 – 1.99) ((0.83 – 1.99))

It follows from the results given in Table 14.1 that  $\theta_{23}$  is close to, but can be different from,  $\pi/4$ ,  $\theta_{12} \cong \pi/5.4$  and that  $\theta_{13} \cong \pi/20$ . Correspondingly, the pattern of neutrino mixing is drastically different from the pattern of quark mixing.

Note also that  $\Delta m_{21}^2$ ,  $\sin^2 \theta_{12}$ ,  $|\Delta m_{31(32)}^2|$ ,  $\sin^2 \theta_{23}$  and  $\sin^2 \theta_{13}$  are determined from the data with a  $1\sigma$  uncertainty ( $= 1/6$  of the  $3\sigma$  range) of approximately 2.3%, 5.8%, 1.7%, 9.0% and 4.8%, respectively.

The existing SK atmospheric neutrino, K2K, MINOS, T2K and NO $\nu$ A data do not allow to determine the sign of  $\Delta m_{31(32)}^2$ . Maximal solar neutrino mixing, i.e.,  $\theta_{12} = \pi/4$ , is ruled out at more than  $6\sigma$  by the data. Correspondingly, one has  $\cos 2\theta_{12} \geq 0.29$  (at 99.73% CL).

Apart from the hint that the Dirac phase  $\delta \cong 3\pi/2$ , no other experimental information on the Dirac and Majorana CPV phases in the neutrino mixing matrix is available at present. Thus, the status of CP symmetry in the lepton sector is essentially unknown. With  $\theta_{13} \cong 0.15 \neq 0$ , the Dirac phase  $\delta$  can generate CP violating effects in neutrino oscillations [48,61,62], i.e., a difference between the probabilities of the  $\nu_l \rightarrow \nu_{l'}$  and  $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$  oscillations,  $l \neq l' = e, \mu, \tau$ . The magnitude of CP violation in  $\nu_l \rightarrow \nu_{l'}$  and  $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$  oscillations,  $l \neq l' = e, \mu, \tau$ , is determined by [63] the rephasing invariant  $J_{CP}$ , associated with the Dirac CPV phase in  $U$ :

$$J_{CP} = \text{Im} \left( U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^* \right). \quad (14.9)$$

It is analogous to the rephasing invariant associated with the Dirac CPV phase in the CKM quark mixing matrix [64]. In the “standard” parametrization of the neutrino mixing matrix (Eq. (14.6)),  $J_{CP}$  has the form:

$$J_{CP} \equiv \text{Im} (U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*) = \frac{1}{8} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \sin \delta. \quad (14.10)$$



Thus, given the fact that  $\sin 2\theta_{12}$ ,  $\sin 2\theta_{23}$  and  $\sin 2\theta_{13}$  have been determined experimentally with a relatively good precision, the size of CP violation effects in neutrino oscillations depends essentially only on the magnitude of the currently not well determined value of the Dirac phase  $\delta$ . The current data implies  $J_{CP} \lesssim 0.035 \sin \delta$ , where we have used the  $3\sigma$  ranges of  $\sin^2 \theta_{12}$ ,  $\sin^2 \theta_{23}$  and  $\sin^2 \theta_{13}$  given in Table 14.1. For the best fit values of  $\sin^2 \theta_{12}$ ,  $\sin^2 \theta_{23}$ ,  $\sin^2 \theta_{13}$  and  $\delta$  we find in the case of  $\Delta m_{31(2)}^2 > 0$  ( $\Delta m_{31(2)}^2 < 0$ ):  $J_{CP} \cong 0.0327 \sin \delta \cong -0.0291$  ( $J_{CP} \cong 0.0327 \sin \delta \cong -0.0276$ ). Thus, if the indication that  $\delta \cong 3\pi/2$  is confirmed by future more precise data, the CP violation effects in neutrino oscillations would be relatively large.

If the neutrinos with definite masses  $\nu_i$ ,  $i = 1, 2, 3$ , are Majorana particles, the 3-neutrino mixing matrix contains two additional Majorana CPV phases [48]. However, the flavour neutrino oscillation probabilities  $P(\nu_l \rightarrow \nu_{l'})$  and  $P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$ ,  $l, l' = e, \mu, \tau$ , do not depend on the Majorana phases [48,65]. The Majorana phases can play important role, *e.g.*, in  $|\Delta L| = 2$  processes like neutrinoless double beta  $((\beta\beta)_{0\nu})$  decay  $(A, Z) \rightarrow (A, Z+2) + e^- + e^-$ ,  $L$  being the total lepton charge, in which the Majorana nature of massive neutrinos  $\nu_i$  manifests itself (see, *e.g.*, Refs. [46,66]). Our interest in the CPV phases present in the neutrino mixing matrix is stimulated also by the intriguing possibility that the Dirac phase and/or the Majorana phases in  $U_{PMNS}$  can provide the CP violation necessary for the generation of the observed baryon asymmetry of the Universe [67,68].

As we have indicated, the existing data do not allow one to determine the sign of  $\Delta m_{31(2)}^2$ . In the case of 3-neutrino mixing, the two possible signs of  $\Delta m_{31(2)}^2$  correspond to two types of neutrino mass spectrum. In the widely used conventions of numbering the neutrinos with definite mass in the two cases, the two spectra read:

– *i) spectrum with normal ordering (NO):*

$$m_1 < m_2 < m_3, \quad \Delta m_{31}^2 = \Delta m_A^2 > 0, \\ \Delta m_{21}^2 \equiv \Delta m_\odot^2 > 0, \quad m_{2(3)} = (m_1^2 + \Delta m_{21(31)}^2)^{1/2}. \quad (14.11)$$

– *ii) spectrum with inverted ordering (IO):*

$$m_3 < m_1 < m_2, \quad \Delta m_{32}^2 = \Delta m_A^2 < 0, \quad \Delta m_{21}^2 \equiv \Delta m_\odot^2 > 0, \\ m_2 = (m_3^2 + \Delta m_{23}^2)^{1/2}, \quad m_1 = (m_3^2 + \Delta m_{23}^2 - \Delta m_{21}^2)^{1/2}. \quad (14.12)$$

Depending on the values of the lightest neutrino mass [69],  $\min(m_j)$ , the neutrino mass spectrum can also be:

– *Normal Hierarchical (NH):*

$$m_1 \ll m_2 < m_3, \quad m_2 \cong (\Delta m_{21}^2)^{1/2} \cong 0.0086 \text{ eV}, \\ m_3 \cong |\Delta m_{31}^2|^{1/2} \cong 0.0504 \text{ eV}; \text{ or} \quad (14.13)$$

– *Inverted Hierarchical (IH):*

$$m_3 \ll m_1 < m_2, \quad m_{1,2} \cong |\Delta m_{32}^2|^{1/2} \cong 0.0500 \text{ eV}; \text{ or} \quad (14.14)$$

– *Quasi-Degenerate (QD):*

$$m_1 \cong m_2 \cong m_3 \cong m_0, \quad m_j^2 \gg |\Delta m_{31(32)}^2|, \quad m_0 \gtrsim 0.10 \text{ eV}. \quad (14.15)$$

Sometimes the determination of the neutrino mass spectrum is referred to in the literature on the subject as determination of “neutrino mass hierarchy”. However, as we have seen, the neutrino mass spectrum might not be hierarchical. Therefore, determination of “neutrino mass ordering” is a more precise expression and we are going to use this expression in the present review article.

Eq. (14.11) and Eq. (14.12) suggest that, for consistency, the data on the larger neutrino mass squared difference, obtained in 3-neutrino oscillation analyses, should be presented i) either on the value of  $\Delta m_{32}^2$  in the case of NO spectrum and on  $\Delta m_{31}^2$  for IO spectrum, or ii) on the value of  $\Delta m_{31}^2$  for the NO spectrum and on  $\Delta m_{32}^2$  for IO spectrum. It would be preferable that all experimental groups which provide data on the larger neutrino mass squared difference, choose one of the indicated two possibilities to present their data - this

will make straightforward the comparison of the results obtained in different experiments.

All types of neutrino mass spectrum, discussed above, are compatible with the existing constraints on the absolute scale of neutrino masses  $m_j$ . Information about the latter can be obtained, *e.g.*, by measuring the spectrum of electrons near the end point in  $^3\text{H}$   $\beta$ -decay experiments [70–74] and from cosmological and astrophysical data. The most stringent upper bounds on the  $\bar{\nu}_e$  mass were obtained in the Troitzk [74,71] experiment:

$$m_{\bar{\nu}_e} < 2.05 \text{ eV} \quad \text{at } 95\% \text{ CL}. \quad (14.16)$$

Similar result was obtained in the Mainz experiment [72]:  $m_{\bar{\nu}_e} < 2.3 \text{ eV}$  at 95% CL. We have  $m_{\bar{\nu}_e} \cong m_{1,2,3}$  in the case of QD spectrum. The KATRIN experiment [73] is planned to reach sensitivity of  $m_{\bar{\nu}_e} \sim 0.20 \text{ eV}$ , *i.e.*, it will probe the region of the QD spectrum.

The Cosmic Microwave Background (CMB) data of the WMAP experiment, combined with supernovae data and data on galaxy clustering can be used to obtain an upper limit on the sum of neutrinos masses (see review on Cosmological Parameters [75] and, *e.g.*, Ref. 76). Depending on the model complexity and the input data used one obtains [76]:  $\sum_j m_j \lesssim (0.3 - 1.3) \text{ eV}$ , 95% CL.

In March of 2013 the Planck Collaboration published their first constraints on  $\sum_j m_j$  [77]. These constraints were updated in 2015 in [78]. Assuming the existence of three light massive neutrinos and the validity of the  $\Lambda$  CDM (Cold Dark Matter) model, and using their data on the CMB temperature power spectrum anisotropies, polarization, on gravitational lensing effects and the low  $l$  CMB polarization spectrum data (the “low P” data), the Planck Collaboration reported the following updated upper limit on the sum of the neutrino masses [78]:  $\sum_j m_j < 0.57 \text{ eV}$ , 95% CL. Adding supernovae (light-curve) data and data on the Baryon Acoustic Oscillations (BAO) lowers the limit to [78]:

$$\sum_j m_j < 0.23 \text{ eV}, \quad 95\% \text{ CL}.$$

It follows from these data that neutrino masses are much smaller than the masses of charged leptons and quarks. If we take as an indicative upper limit  $m_j \lesssim 0.5 \text{ eV}$ , we have  $m_j/m_{l,q} \lesssim 10^{-6}$ ,  $l = e, \mu, \tau$ ,  $q = d, s, b, u, c, t$ . It is natural to suppose that the remarkable smallness of neutrino masses is related to the existence of a new fundamental mass scale in particle physics, and thus to new physics beyond that predicted by the Standard Model.

### 14.3. Future progress

After the spectacular experimental progress made in the studies of neutrino oscillations, further understanding of the pattern of neutrino masses and neutrino mixing, of their origins and of the status of CP symmetry in the lepton sector requires an extensive and challenging program of research. The main goals of such a research program include:

- Determining the nature - Dirac or Majorana, of massive neutrinos  $\nu_j$ . This is of fundamental importance for making progress in our understanding of the origin of neutrino masses and mixing and of the symmetries governing the lepton sector of particle interactions.
- Determination of the sign of  $\Delta m_{31}^2$  (or  $\Delta m_{32}^2$ ), *i.e.*, the “neutrino mass ordering”, or of the type of spectrum neutrino masses obey.
- Determining, or obtaining significant constraints on, the absolute scale of neutrino masses. This, in particular, would help obtain information about the detailed structure (hierarchical, quasidegenerate, etc.) of the neutrino mass spectrum.
- Determining the status of CP symmetry in the lepton sector.
- High precision measurement of  $\theta_{13}$ ,  $\Delta m_{21}^2$ ,  $\theta_{12}$ ,  $|\Delta m_{31}^2|$  and  $\theta_{23}$ .
- Understanding at a fundamental level the mechanism giving rise to neutrino masses and mixing and to  $L_1$ -non-conservation. This includes understanding the origin of the patterns of  $\nu$ -mixing

and  $\nu$ -masses suggested by the data. Are the observed patterns of  $\nu$ -mixing and of  $\Delta m_{21,31}^2$  related to the existence of a new fundamental symmetry of particle interactions? Is there any relation between quark mixing and neutrino mixing, *e.g.*, does the relation  $\theta_{12} + \theta_c = \pi/4$ , where  $\theta_c$  is the Cabibbo angle, hold? What is the physical origin of CP violation phases in the neutrino mixing matrix  $U$ ? Is there any relation (correlation) between the (values of) CP violation phases and mixing angles in  $U$ ? Progress in the theory of neutrino mixing might also lead to a better understanding of the mechanism of generation of baryon asymmetry of the Universe.

The high precision measurement of the value of  $\sin^2 2\theta_{13}$  from the Daya Bay experiment and the subsequent results on  $\theta_{13}$  obtained by the RENO, Double Chooz and T2K collaborations (see Section 1.11), have far reaching implications. The measured relatively large value of  $\theta_{13}$  opened up the possibilities, in particular,

- i) for searching for CP violation effects in neutrino oscillation experiments with high intensity accelerator neutrino beams, like T2K, NO $\nu$ A, etc. The sensitivities of T2K and NO $\nu$ A on CP violation in neutrino oscillations are discussed in, *e.g.*, Refs. [79,80].
- ii) for determining the sign of  $\Delta m_{32}^2$ , and thus the type of neutrino mass spectrum (“neutrino mass ordering”) in the long baseline neutrino oscillation experiments at accelerators (NO $\nu$ A, etc.), in the experiments studying the oscillations of atmospheric neutrinos (PINGU, ORCA), as well as in experiments with reactor antineutrinos [81] (for a review see, *e.g.*, Ref. 82).

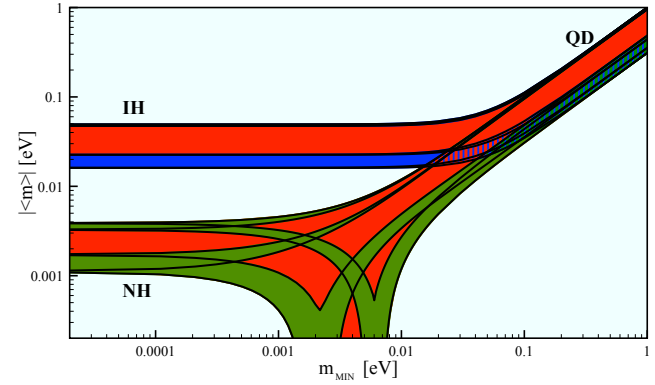
There are also long term plans extending beyond 2025 for searches for CP violation and neutrino mass spectrum (ordering) determination in long baseline neutrino oscillation experiments with accelerator neutrino beams (see, *e.g.*, Refs. [80,83,84,85]). The successful realization of this research program would be a formidable task and would require many years of extraordinary experimental efforts aided by intensive theoretical investigations and remarkable investments.

Before reviewing in detail i) the different neutrino sources and the specific characteristics of the corresponding neutrino fluxes, which have been and are being used in neutrino oscillation experiments (Section 14.6), ii) the theory and phenomenology of neutrino oscillations (Sections 14.7 and 14.8), and iii) the compelling experimental evidences of neutrino oscillations and, more generally, the results obtained in the neutrino oscillation experiments (Sections 14.9 and 14.13), we would like to discuss briefly the problem of determination of the nature - Dirac or Majorana - of massive neutrinos as well as the (type I) seesaw mechanism of neutrino mass generation.

#### 14.4. The nature of massive neutrinos

The experiments studying flavour neutrino oscillations cannot provide information on the nature - Dirac or Majorana, of massive neutrinos [48,65]. Establishing whether the neutrinos with definite mass  $\nu_j$  are Dirac fermions possessing distinct antiparticles, or Majorana fermions, *i.e.*, spin 1/2 particles that are identical with their antiparticles, is of fundamental importance for understanding the origin of  $\nu$ -masses and mixing and the underlying symmetries of particle interactions (see, *e.g.*, Ref. 86). The neutrinos with definite mass  $\nu_j$  will be Dirac fermions if the particle interactions conserve some additive lepton number, *e.g.*, the total lepton charge  $L = L_e + L_\mu + L_\tau$ . If no lepton charge is conserved,  $\nu_j$  will be Majorana fermions (see, *e.g.*, Ref. 46). The massive neutrinos are predicted to be of Majorana nature by the see-saw mechanism of neutrino mass generation [3]. The observed patterns of neutrino mixing and of neutrino mass squared differences can be related to Majorana massive neutrinos and the existence of an approximate flavour symmetry in the lepton sector (see, *e.g.*, Ref. 87). Determining the nature of massive neutrinos  $\nu_j$  is one of the fundamental and most challenging problems in the future studies of neutrino mixing.

The Majorana nature of massive neutrinos  $\nu_j$  manifests itself in the existence of processes in which the total lepton charge  $L$  changes by two units:  $K^+ \rightarrow \pi^- + \mu^+ + \mu^+$ ,  $\mu^- + (A, Z) \rightarrow \mu^+ + (A, Z - 2)$ , *etc.* Extensive studies have shown that the only feasible experiments having the potential of establishing that the massive neutrinos are Majorana



**Figure 14.1:** The effective Majorana mass  $\langle m \rangle$  (including a  $2\sigma$  uncertainty), as a function of  $\min(m_j)$ . The figure is obtained using the best fit values and the  $2\sigma$  ranges of allowed values of  $\Delta m_{21}^2$ ,  $\sin^2 \theta_{12}$ , and  $|\Delta m_{31}^2| \cong |\Delta m_{32}^2|$  from Ref. 60. The phases  $\alpha_{21,31}$  are varied in the interval  $[0, \pi]$ . The predictions for the NH, IH and QD spectra are indicated. The red regions correspond to at least one of the phases  $\alpha_{21,31}$  and  $(\alpha_{31} - \alpha_{21})$  having a CP violating value, while the blue and green areas correspond to  $\alpha_{21,31}$  possessing CP conserving values. (Update by S. Pascoli of a figure from the one before the last article quoted in Ref. 91.)

particles are at present the experiments searching for  $(\beta\beta)_{0\nu}$ -decay:  $(A, Z) \rightarrow (A, Z + 2) + e^- + e^-$  (see, *e.g.*, Ref. 88). The observation of  $(\beta\beta)_{0\nu}$ -decay and the measurement of the corresponding half-life with sufficient accuracy, would not only be a proof that the total lepton charge is not conserved, but might also provide unique information on the i) type of neutrino mass spectrum (see, *e.g.*, Ref. 89), ii) Majorana phases in  $U$  [66,90] and iii) the absolute scale of neutrino masses (for details see Ref. 88 to Ref. 91 and references quoted therein).

Under the assumptions of 3- $\nu$  mixing, of massive neutrinos  $\nu_j$  being Majorana particles, and of  $(\beta\beta)_{0\nu}$ -decay generated only by the (V-A) charged current weak interaction via the exchange of the three Majorana neutrinos  $\nu_j$  having masses  $m_j \lesssim$  few MeV, the  $(\beta\beta)_{0\nu}$ -decay amplitude has the form (see, *e.g.*, Ref. 46 and Ref. 88):  $A(\beta\beta)_{0\nu} \cong \langle m \rangle M$ , where  $M$  is the corresponding nuclear matrix element which does not depend on the neutrino mixing parameters, and

$$\begin{aligned} \langle m \rangle &= \left| m_1 U_{e1}^2 + m_2 U_{e2}^2 + m_3 U_{e3}^2 \right| \\ &= \left| \left( m_1 c_{12}^2 + m_2 s_{12}^2 e^{i\alpha_{21}} \right) c_{13}^2 + m_3 s_{13}^2 e^{i(\alpha_{31} - 2\delta)} \right|, \end{aligned} \quad (14.17)$$

is the effective Majorana mass in  $(\beta\beta)_{0\nu}$ -decay. In the case of CP-invariance one has [50],  $\eta_{21} \equiv e^{i\alpha_{21}} = \pm 1$ ,  $\eta_{31} \equiv e^{i\alpha_{31}} = \pm 1$ ,  $e^{-i2\delta} = 1$ . The three neutrino masses  $m_{1,2,3}$  can be expressed in terms of the two measured  $\Delta m_{jk}^2$  and, *e.g.*,  $\min(m_j)$ . Thus, given the neutrino oscillation parameters  $\Delta m_{21}^2$ ,  $\sin^2 \theta_{12}$ ,  $\Delta m_{31}^2$  and  $\sin^2 \theta_{13}$ ,  $\langle m \rangle$  is a function of the lightest neutrino mass  $\min(m_j)$ , the Majorana (and Dirac) CP violation phases in  $U$  and of the type of neutrino mass spectrum. In the case of NH, IH and QD spectrum we have (see, *e.g.*, Ref. 66 and Ref. 91):

$$\langle m \rangle \cong \left| \sqrt{\Delta m_{21}^2 s_{12}^2 c_{13}^2} + \sqrt{\Delta m_{31}^2 s_{13}^2} e^{i(\alpha_{31} - \alpha_{21} - 2\delta)} \right|, \quad \text{NH}, \quad (14.18)$$

$$\langle m \rangle \cong \tilde{m} \left( 1 - \sin^2 2\theta_{12} \sin^2 \frac{\alpha_{21}}{2} \right)^{\frac{1}{2}}, \quad \text{IH (IO) and QD}, \quad (14.19)$$

where  $\tilde{m} \equiv \sqrt{\Delta m_{23}^2 + m_3^2}$  and  $\tilde{m} \equiv m_0$  for IH (IO) and QD spectrum, respectively. In Eq. (14.19) we have exploited the fact that  $\sin^2 \theta_{13} \ll \cos 2\theta_{12}$ . The CP conserving values of the Majorana phase  $\alpha_{21}$  and the Majorana-Dirac phase difference  $(\alpha_{31} - \alpha_{21} - 2\delta)$  determine the intervals of possible values of  $\langle m \rangle$ , corresponding to the different types of neutrino mass spectrum. Using the  $3\sigma$  ranges of the allowed

values of the neutrino oscillation parameters from Table 14.1 one finds that: i)  $0.81 \times 10^{-3} \text{ eV} \lesssim |<m>| \lesssim 4.43 \times 10^{-3} \text{ eV}$  in the case of NH spectrum; ii)  $\sqrt{\Delta m_{23}^2} \cos 2\theta_{12} c_{13}^2 \lesssim |<m>| \lesssim \sqrt{\Delta m_{23}^2} c_{13}^2$ , or  $1.3 \times 10^{-2} \text{ eV} \lesssim |<m>| \lesssim 5.0 \times 10^{-2} \text{ eV}$  in the case of IH spectrum;  $1.3 \times 10^{-2} \text{ eV} \lesssim |<m>| \lesssim 5.0 \times 10^{-2} \text{ eV}$  in the case of IH spectrum; iii)  $m_0 \cos 2\theta_{12} \lesssim |<m>| \lesssim m_0$ , or  $2.9 \times 10^{-2} \text{ eV} \lesssim |<m>| \lesssim m_0 \text{ eV}$ ,  $m_0 \gtrsim 0.10 \text{ eV}$ , in the case of QD spectrum. The difference in the ranges of  $|<m>|$  in the cases of NH, IH and QD spectrum opens up the possibility to get information about the type of neutrino mass spectrum from a measurement of  $|<m>|$  [89]. The predicted  $(\beta\beta)_{0\nu}$ -decay effective Majorana mass  $|<m>|$  as a function of the lightest neutrino mass  $\min(m_j)$  is shown in Fig. 14.1.

The  $(\beta\beta)_{0\nu}$ -decay can be generated, in principle, by a  $\Delta L = 2$  mechanism other than the light Majorana neutrino exchange considered here, or by a combination of mechanisms one of which is the light Majorana neutrino exchange (for a discussion of different mechanisms which can trigger  $(\beta\beta)_{0\nu}$ -decay, see, *e.g.*, Refs. [92,93] and the articles quoted therein). If the  $(\beta\beta)_{0\nu}$ -decay will be observed, it will be of fundamental importance to determine which mechanism (or mechanisms) is (are) inducing the decay. The discussion of the problem of determining the mechanisms which possibly are operative in  $(\beta\beta)_{0\nu}$ -decay, including the case when more than one mechanism is involved, is out of the scope of the present article. This problem has been investigated in detail in, *e.g.*, Refs. [93,94] and we refer the reader to these articles and the articles quoted therein.

#### 14.5. The see-saw mechanism and the baryon asymmetry of the Universe

A natural explanation of the smallness of neutrino masses is provided by the (type I) see-saw mechanism of neutrino mass generation [3]. An integral part of this rather simple mechanism [95] are the RH neutrinos  $\nu_{lR}$  (RH neutrino fields  $\nu_{lR}(x)$ ). The latter are assumed to possess a Majorana mass term as well as Yukawa type coupling  $\mathcal{L}_Y(x)$  with the Standard Model lepton and Higgs doublets,  $\psi_{lL}(x)$  and  $\Phi(x)$ , respectively,  $(\psi_{lL}(x))^T = (\nu_{lL}^T(x) \quad l_L^T(x))$ ,  $l = e, \mu, \tau$ ,  $(\Phi(x))^T = (\Phi^{(0)}(x) \quad \Phi^{(-)}(x))$ . In the basis in which the Majorana mass matrix of RH neutrinos is diagonal, we have:

$$\mathcal{L}_{Y,M}(x) = \left( \lambda_{il} \overline{N_i}(x) \Phi^\dagger(x) \psi_{lL}(x) + \text{h.c.} \right) - \frac{1}{2} M_i \overline{N_i}(x) N_i(x), \quad (14.20)$$

where  $\lambda_{il}$  is the matrix of neutrino Yukawa couplings and  $N_i$  ( $N_i(x)$ ) is the heavy RH Majorana neutrino (field) possessing a mass  $M_i > 0$ . When the electroweak symmetry is broken spontaneously, the neutrino Yukawa coupling generates a Dirac mass term:  $m_{il}^D \overline{N_i}(x) \nu_{lL}(x) + \text{h.c.}$ , with  $m^D = v\lambda$ ,  $v = 174 \text{ GeV}$  being the Higgs doublet v.e.v. In the case when the elements of  $m^D$  are much smaller than  $M_k$ ,  $|m_{il}^D| \ll M_k$ ,  $i, k = 1, 2, 3$ ,  $l = e, \mu, \tau$ , the interplay between the Dirac mass term and the mass term of the heavy (RH) Majorana neutrinos  $N_i$  generates an effective Majorana mass (term) for the LH flavour neutrinos [3]:

$$m_{ll'}^{LL} \cong -(m^D)_{l'j}^T M_j^{-1} m_{jl}^D = -v^2 (\lambda)_{l'j}^T M_j^{-1} \lambda_{jl}. \quad (14.21)$$

In grand unified theories,  $m^D$  is typically of the order of the charged fermion masses. In  $SO(10)$  theories, for instance,  $m^D$  coincides with the up-quark mass matrix. Taking indicatively  $m^{LL} \sim 0.1 \text{ eV}$ ,  $m^D \sim 100 \text{ GeV}$ , one finds  $M \sim 10^{14} \text{ GeV}$ , which is close to the scale of unification of the electroweak and strong interactions,  $M_{GUT} \cong 2 \times 10^{16} \text{ GeV}$ . In GUT theories with RH neutrinos one finds that indeed the heavy Majorana neutrinos  $N_j$  naturally obtain masses which are by few to several orders of magnitude smaller than  $M_{GUT}$ . Thus, the enormous disparity between the neutrino and charged fermion masses is explained in this approach by the huge difference between effectively the electroweak symmetry breaking scale and  $M_{GUT}$ .

An additional attractive feature of the see-saw scenario is that the generation and smallness of neutrino masses is related via the leptogenesis mechanism [2] to the generation of the baryon

asymmetry of the Universe. The Yukawa coupling in Eq. (14.20), in general, is not CP conserving. Due to this CP-nonconserving coupling the heavy Majorana neutrinos undergo, *e.g.*, the decays  $N_j \rightarrow l^+ + \Phi^{(-)}$ ,  $N_j \rightarrow l^- + \Phi^{(+)}$ , which have different rates:  $\Gamma(N_j \rightarrow l^+ + \Phi^{(-)}) \neq \Gamma(N_j \rightarrow l^- + \Phi^{(+)})$ . When these decays occur in the Early Universe at temperatures somewhat below the mass of, say,  $N_1$ , so that the latter are out of equilibrium with the rest of the particles present at that epoch, CP violating asymmetries in the individual lepton charges  $L_l$ , and in the total lepton charge  $L$ , of the Universe are generated. These lepton asymmetries are converted into a baryon asymmetry by  $(B - L)$  conserving, but  $(B + L)$  violating, sphaleron processes, which exist in the Standard Model and are effective at temperatures  $T \sim (100 - 10^{12}) \text{ GeV}$ . If the heavy neutrinos  $N_j$  have hierarchical spectrum,  $M_1 \ll M_2 \ll M_3$ , the observed baryon asymmetry can be reproduced provided the mass of the lightest one satisfies  $M_1 \gtrsim 10^9 \text{ GeV}$  [96]. Thus, in this scenario, the neutrino masses and mixing and the baryon asymmetry have the same origin - the neutrino Yukawa couplings and the existence of (at least two) heavy Majorana neutrinos. Moreover, quantitative studies [67] based on advances in leptogenesis theory [97] have shown that the CP violation, necessary in leptogenesis for the generation of the observed baryon asymmetry of the Universe, can be provided exclusively by the Dirac and/or Majorana phases in the neutrino mixing matrix  $U$ . This implies, in particular, that if the CP symmetry is established not to hold in the lepton sector due to  $U$ , at least some fraction (if not all) of the observed baryon asymmetry might be due to the Dirac and/or Majorana CP violation present in the neutrino mixing. More specifically, the necessary condition that the requisite CP violation for a successful leptogenesis with hierarchical in mass heavy Majorana neutrinos is due *entirely* to the Dirac CPV phase in  $U$  reads [68]:  $|\sin \theta_{13} \sin \delta| \gtrsim 0.09$ . This condition is comfortably compatible with the measured value of  $\sin \theta_{13} \cong 0.15$  and the hint that  $\delta \cong 3\pi/2$ , found in the global analyses of the neutrino oscillation data.

#### 14.6. Neutrino sources

In the experimental part of this review (Sections 14.9 - 14.13), we mainly discuss neutrino oscillation experiments using neutrinos or antineutrinos produced by the Sun, cosmic-ray interactions in the air, proton accelerators, and nuclear reactors. We call neutrinos from these sources as solar neutrinos, atmospheric neutrinos, accelerator neutrinos, and reactor (anti)neutrinos. Neutrinos (and/or antineutrinos) from each of these sources have very different properties, *e.g.*, energy spectra, flavour components, and directional distributions, at production. In the literature, neutrino flavour conversion of neutrinos from gravitationally collapsed supernova explosions (supernova neutrinos) is also discussed, but this topic is out of the scope of the present review.

Solar neutrinos and atmospheric neutrinos are naturally produced neutrinos; their fluxes as well as the distance between the (point or distributed) neutrino source and the detector cannot be controlled artificially. While the atmospheric neutrino flux involves  $\nu_\mu$ ,  $\bar{\nu}_\mu$ ,  $\nu_e$ , and  $\bar{\nu}_e$  components at production, solar neutrinos are produced as pure electron neutrinos due to thermo-nuclear fusion reactions of four protons, producing a helium nucleus. For atmospheric neutrinos with energy  $\gtrsim 1 \text{ GeV}$ , which undergo charged-current interactions in the detector, directional correlation of the charged lepton with the parent neutrino gives the way to know, within the resolution, the distance traveled by the neutrino between the production and detection.

Accelerator neutrinos and reactor (anti)neutrinos are man-made neutrinos. In principle, it is possible to choose the distance between the neutrino source and the detector arbitrarily. Accelerator neutrinos used for neutrino oscillation experiments so far have been produced by the decay of secondary mesons (pions and kaons) produced by the collision of a primary proton beam with a nuclear target. A dominant component of the accelerator neutrino flux is  $\nu_\mu$  or  $\bar{\nu}_\mu$ , depending on the secondary meson's sign selection, but a wrong-sign muon neutrino component as well as  $\nu_e$  and  $\bar{\nu}_e$  components are also present. The fluxes of the accelerator neutrinos depend on a number of factors, *e.g.*, energy and intensity of the primary proton beam, material and geometry of the target, selection of the momentum and charge of

the secondary mesons that are focused, and production angle of the secondary mesons with respect to the primary beam. In other words, it is possible to control the peak energy, energy spread, and dominant neutrino flavour, of the neutrino beam.

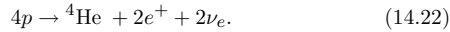
From the nuclear reactor, almost pure electron antineutrinos are produced by  $\beta$ -decays of fission products of the nuclear fuel. However, experimental groups cannot control the normalization and spectrum of the  $\bar{\nu}_e$  flux from commercial nuclear reactors. They are dependent on the initial fuel composition and history of the nuclear fuel burnup. These data are provided by the power plant companies.

For neutrino oscillation experiments, knowledge of the flux of each neutrino and antineutrino flavour at production is needed for planning and designing the experiment, analyzing the data, and estimating systematic errors. Basically, for all neutrino sources, flux models are constructed and validation is made by comparing various experimentally observed quantities with the model predictions. Many of the modern accelerator long baseline and reactor neutrino oscillation experiments employ a two- or multi-detector configuration. In the accelerator long baseline experiment, a “near” detector measures non-oscillated neutrino flux. In the two- or multi-baseline reactor experiments, even a near detector measures the neutrino flux with oscillations developed to some extent. However, comparing the quantities measured with different baselines, it is possible to validate the reactor flux model and measure the oscillation parameters at the same time, or to make an analysis with minimal dependence on flux models.

#### 14.6.1. Standard solar model predictions of the solar neutrino fluxes :

Observation of solar neutrinos directly addresses the theory of stellar structure and evolution, which is the basis of the standard solar model (SSM). The Sun, as a well-defined neutrino source, also provides an important opportunities to investigate neutrino oscillations including matter effects, because of the wide range of matter density and the great distance from the Sun to the Earth.

The solar neutrinos are produced by some of the fusion reactions in the pp chain or CNO cycle. The combined effect of these reactions is written as



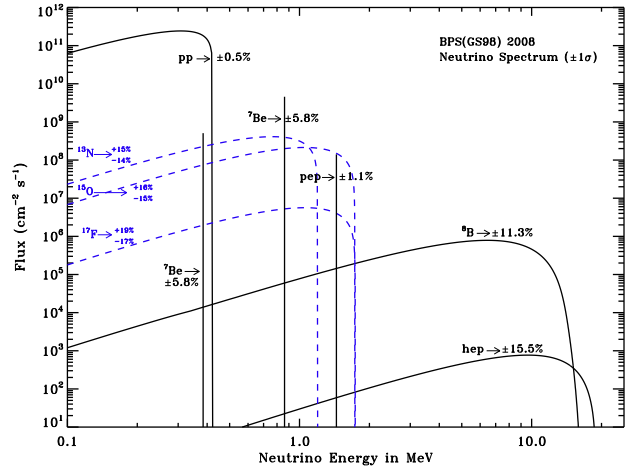
**Table 14.2:** Neutrino-producing reactions in the Sun (first column) and their abbreviations (second column). The neutrino fluxes predicted by the BPS08(GS) model [100] are listed in the third column.

Reaction	Abbr.	Flux ( $\text{cm}^{-2} \text{s}^{-1}$ )
$pp \rightarrow de^+ \nu$	$pp$	$5.97(1 \pm 0.006) \times 10^{10}$
$pe^-p \rightarrow d\nu$	$pep$	$1.41(1 \pm 0.011) \times 10^8$
${}^3\text{He} p \rightarrow {}^4\text{He} e^+ \nu$	$hep$	$7.90(1 \pm 0.15) \times 10^3$
${}^7\text{Be} e^- \rightarrow {}^7\text{Li} \nu + (\gamma)$	${}^7\text{Be}$	$5.07(1 \pm 0.06) \times 10^9$
${}^8\text{B} \rightarrow {}^8\text{Be}^* e^+ \nu$	${}^8\text{B}$	$5.94(1 \pm 0.11) \times 10^6$
${}^{13}\text{N} \rightarrow {}^{13}\text{C} e^+ \nu$	${}^{13}\text{N}$	$2.88(1 \pm 0.15) \times 10^8$
${}^{15}\text{O} \rightarrow {}^{15}\text{N} e^+ \nu$	${}^{15}\text{O}$	$2.15(1^{+0.17}_{-0.16}) \times 10^8$
${}^{17}\text{F} \rightarrow {}^{17}\text{O} e^+ \nu$	${}^{17}\text{F}$	$5.82(1^{+0.19}_{-0.17}) \times 10^6$

Positrons annihilate with electrons. Therefore, when considering the solar thermal energy generation, a relevant expression is

$$4p + 2e^- \rightarrow {}^4\text{He} + 2\nu_e + 26.73 \text{ MeV} - E_\nu, \quad (14.23)$$

where  $E_\nu$  represents the energy taken away by neutrinos, the average value being  $\langle E_\nu \rangle \sim 0.6 \text{ MeV}$ . There have been efforts to calculate solar neutrino fluxes from these reactions on the basis of SSM. A variety of input information is needed in the evolutionary calculations. The most elaborate SSM calculations have been developed by Bahcall and his collaborators, who define their SSM as the solar model which is



**Figure 14.2:** The solar neutrino spectrum predicted by the BPS08(GS) standard solar model [100]. The neutrino fluxes are given in units of  $\text{cm}^{-2}\text{s}^{-1}\text{MeV}^{-1}$  for continuous spectra and  $\text{cm}^{-2}\text{s}^{-1}$  for line spectra. The numbers associated with the neutrino sources show theoretical errors of the fluxes. This figure is taken from Aldo Serenelli’s web site, <http://www.mpa-garching.mpg.de/~aldos/>.

constructed with the best available physics and input data. Therefore, their SSM calculations have been rather frequently updated. SSM’s labelled as BS05(OP) [98], BSB06(GS) and BSB06(AGS) [99], and BPS08(GS) and BPS08(AGS) [100] represent some of the relatively recent model calculations. Here, “OP” means that newly calculated radiative opacities from the “Opacity Project” are used. The later models are also calculated with OP opacities. “GS” and “AGS” refer to old and new determinations of solar abundances of heavy elements. There are significant differences between the old, higher heavy element abundances (GS) and the new, lower heavy element abundances (AGS). The models with GS are consistent with helioseismological data, but the models with AGS are not.

The prediction of the BPS08(GS) model for the fluxes from neutrino-producing reactions is given in Table 14.2. Fig. 14.2 shows the solar-neutrino spectra calculated with the BPS08(GS) model. Here we note that in Ref. 101 the authors point out that electron capture on  ${}^{13}\text{N}$ ,  ${}^{15}\text{O}$ , and  ${}^{17}\text{F}$  produces line spectra of neutrinos, which have not been considered in the SSM calculations quoted above.

In 2011, a new SSM calculations [102] have been presented by A.M. Serenelli, W.C. Haxton, and C. Peña-Garay, by adopting the newly analyzed nuclear fusion cross sections. Their high metallicity SSM is labelled as SHP11(GS). For the same solar abundances as used in Ref. 98 and Ref. 99, the most significant change is a decrease of  ${}^8\text{B}$  flux by  $\sim 5\%$ .

#### 14.6.2. Atmospheric neutrino fluxes :

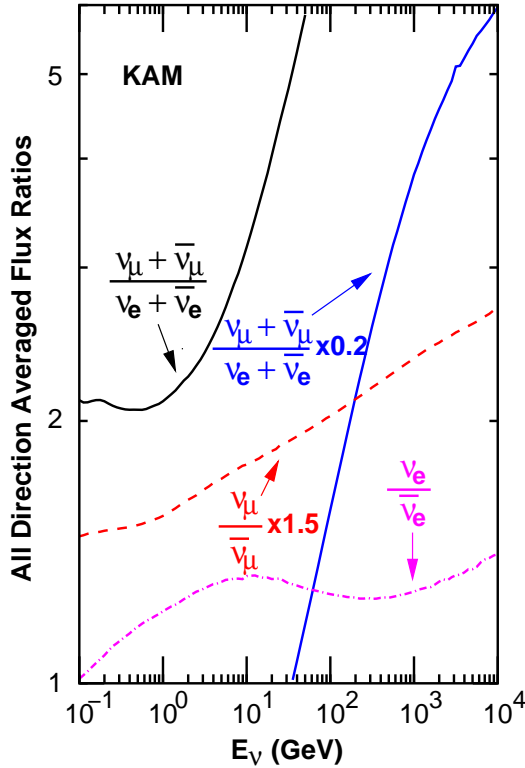
Atmospheric neutrinos are produced by the decay of  $\pi$  and  $K$  mesons produced in the nuclear interactions of the primary component of cosmic rays in the atmosphere ((A) in Table 14.3). The primary cosmic ray components above 2 GeV/nucleon are protons ( $\sim 95\%$ ), helium nuclei ( $\sim 4.5\%$ ), and heavier nuclei. For neutrino producing hadronic interactions, a nucleus can be simply regarded as a sum of individual nucleons at high energies. Pions are dominantly produced in these interactions, and they predominantly decay according to (B1) in Table 14.3, followed by muon decay (E) in Table 14.3. The interactions in massive underground detectors of atmospheric neutrinos provide a means of studying neutrino oscillations because of the large range of distances traveled by these neutrinos ( $\sim 10$  to  $1.27 \times 10^4 \text{ km}$ ) to reach a detector on Earth and relatively well-understood atmospheric neutrino fluxes.

Calculation of the atmospheric neutrino fluxes requires knowledge of the primary cosmic-ray fluxes and composition, and the hadronic interactions. Atmospheric neutrinos with energy of  $\sim$  a few GeV are mostly produced by primary cosmic rays with energy of  $\sim 100 \text{ GeV}$ .

**Table 14.3:** Reactions and decays relevant to atmospheric neutrino and accelerator neutrino production. The first column shows the index of the reaction or decay, and the second column shows the reaction or decay channel. The third column shows the branching ratio [103]. (For  $K_L$  decay, sum of the branching ratios to charge conjugate modes is shown).

	Reaction/Decay	Branching ratio (%)
(A)	$p(n) + A \rightarrow \pi^\pm X, K^\pm X, K_L X$	
(B1)	$\pi^\pm \rightarrow \mu^\pm + \nu_\mu (\bar{\nu}_\mu)$	99.9877
(B2)	$\rightarrow e^\pm + \nu_e (\bar{\nu}_e)$	0.0123
(C1)	$K^\pm \rightarrow \mu^\pm + \nu_\mu (\bar{\nu}_\mu)$	63.55
(C2)	$\rightarrow \pi^0 + \mu^\pm + \nu_\mu (\bar{\nu}_\mu)$	3.353
(C3)	$\rightarrow \pi^0 + e^\pm + \nu_e (\bar{\nu}_e)$	5.07
(D1)	$K_L \rightarrow \pi^\pm + \mu^\mp + \bar{\nu}_\mu (\nu_\mu)$	27.04
(D2)	$\rightarrow \pi^\pm + e^\mp + \bar{\nu}_e (\nu_e)$	40.55
(E)	$\mu^\pm \rightarrow e^\pm + \bar{\nu}_\mu (\nu_\mu) + \nu_e (\bar{\nu}_e)$	100

For primary cosmic-rays in this energy range, a flux modulation due to the solar activity and the effects of Earth's geomagnetic fields should be taken into account. In particular, the atmospheric neutrino fluxes in the low-energy region depend on the location on the Earth. Detailed calculations of the atmospheric neutrino fluxes are performed by Honda et al. [104,105], Barr et al. [106], and Battistoni et al. [107], with a typical uncertainty of 10 ~ 20%.



**Figure 14.3:** Neutrino flavour ratios calculated with the all-direction and one-year averaged atmospheric neutrino fluxes at Kamioka. This figure is provided by M. Honda, and is a part of Fig. 5 in Ref. [105], where similar plots at the INO site, the South Pole, and the Pyhäsalmi mine are also shown.

From the dominant production mechanism of the atmospheric neutrinos, we can readily understand some relations that exist between the atmospheric  $\nu_\mu$ ,  $\bar{\nu}_\mu$ ,  $\nu_e$ , and  $\bar{\nu}_e$  fluxes without detailed calculations. For the ratio of the fluxes of  $(\nu_\mu + \bar{\nu}_\mu)$  and  $(\nu_e + \bar{\nu}_e)$  at low energies

( $\lesssim 1$  GeV), where almost all produced muons decay before reaching the ground, we have approximately  $(\nu_\mu + \bar{\nu}_\mu)/(\nu_e + \bar{\nu}_e) \approx 2$ . As the neutrino energy increases, this ratio increases because an increasing fraction of muons do not decay before reaching the ground and being absorbed. We also have  $\nu_\mu/\bar{\nu}_\mu \approx 1$  at low energies. However, as the  $\nu_e/\bar{\nu}_e$  ratio reflects the parent  $\pi^+/\pi^-$  ratio, it is expected to be slightly greater than 1 because the dominance of protons in the primary component of the cosmic rays means a  $\pi^+$  excess in the secondary component. Fig. 14.3 shows these ratios at the Super-Kamiokande site, averaged over all directions and over a year, as a function of neutrino energy.

Another important feature of the atmospheric neutrino fluxes is that the zenith angle distribution for each neutrino type is up-down symmetric above  $\sim 1$  GeV, if there are no neutrino oscillations. As the neutrino energy becomes lower than  $\sim 1$  GeV, however, zenith angle distributions start to show deviations from up-down symmetric shapes due to the geomagnetic effects on primary cosmic rays.

#### 14.6.3. Accelerator neutrino beams :

Conventional method to produce neutrino beams at a high-energy proton accelerator facility is to guide an intense proton beam onto a nuclear target of  $1 \sim 2$  interaction lengths. For a comprehensive description of the accelerator neutrino beams, see Ref. 108. From the  $pA$  collisions, mesons are produced and their decays then produce neutrinos (see Table 14.3). In the high-energy collisions, pions are dominantly produced, with kaons produced at an order of 10% of the pion production rate. Therefore, the dominant component of the accelerator neutrinos is the muon neutrino or muon antineutrino. Mesons decay in the free space called a decay pipe or decay tunnel. This free space is evacuated or filled with helium gas.

To increase the neutrino flux, it is necessary to focus the secondary pions. Modern neutrino oscillation experiments at high-energy accelerators exploit two or three magnetic horns as an approximately point-to-parallel focusing system for this purpose. A magnetic horn is a high-current pulse magnet with toroidal magnetic fields. Therefore, the use of horns also means sign selection of the secondary hadrons that are focused, which in turn means muon neutrino sign selection. Even so, a fraction of wrong sign muon neutrinos contaminate the beam. Also, there is a small  $\nu_e$  and  $\bar{\nu}_e$  contamination from kaon, pion, and muon decay ((C3), (B2), and (E) in Table 14.3). Precise knowledge of  $\nu_e$  and  $\bar{\nu}_e$  components in the neutrino flux is important for the  $\nu_\mu \rightarrow \nu_e$  ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ ) appearance measurement.

With a given neutrino beam line configuration, the expected neutrino fluxes are calculated by using a simulation program tuned to that configuration. Re-interactions of the primary protons in the target and interactions of the secondary particles in the target and in the material outside the target have to be taken into account. An important input is hadron production cross sections from  $pA$  collisions for relevant target materials over wide energy and angular regions. For this purpose, some dedicated experiments such as SPY [109], HARP [110], MIPP [111], and NA61/SHINE [112] have been conducted. The data are fit to specific hadron production models to determine the model parameters.

The predicted neutrino fluxes have to be validated in some way. Modern long baseline neutrino oscillation experiments often have a two-detector configuration, with a near detector to measure an unoscillated neutrino flux immediately after the production. In the single detector experiment, the muon-neutrino flux model is calibrated by using a muon monitor which is located behind the beam dump. Since low-energy muons are absorbed in the beam dump, it is not possible to calibrate the low-energy part of the neutrino spectrum. Even in the two-detector experiments, it should be noted that the near detector does not see the same neutrino flux as the far detector sees, because the neutrino source looks like a line source for the near detector, while it looks as a point source for the far detector.

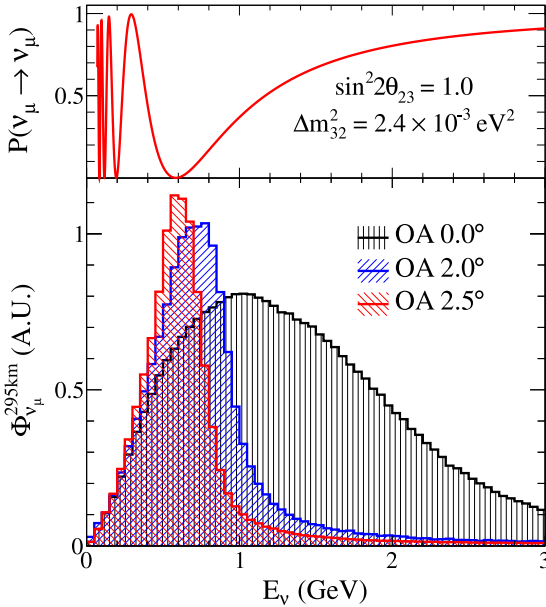
The energy  $E_\nu$  of the neutrino emitted at an angle  $\theta$  with respect to the parent pion direction is given by

$$E_\nu = \frac{m_\pi^2 - m_\mu^2}{2(E_\pi - p_\pi \cos \theta)}, \quad (14.24)$$



where  $E_\pi$  and  $p_\pi$  are the energy and momentum of the parent pion,  $m_\pi$  is the charged pion mass,  $m_\mu$  is the muon mass. Suppose an ideal case that the pions are completely focused in parallel. Then, for  $\theta = 0$ , it can be seen from the above equation that  $E_\nu$  is proportional to  $E_\pi$  for  $E_\pi \gg m_\pi$ . As the secondary pions have a wide energy spectrum, a 0 degree neutrino beam also has a wide spectrum and is called a “wide-band beam”.

For a given angle  $\theta$ , differentiating the above expression with respect to  $E_\pi$ , it can be shown that  $E_\nu$  takes a maximum value  $E_\nu^{\max} = (m_\pi^2 - m_\mu^2)/(2E_\pi^0 \sin^2 \theta)$  at  $E_\pi^0 = m_\pi/\sin \theta$ . Numerical calculations show that a wide range of  $E_\pi$ , in particular that of  $E_\pi \geq E_\pi^0$ , contributes to a narrow range of  $E_\nu \leq E_\nu^{\max}$  [113]. It is expected, therefore, that a narrow neutrino spectrum peaked at around  $E_\nu^{\max}$  can be obtained by the off-axis beam. Fig. 14.4 shows an example of the simulated muon neutrino fluxes at  $\theta = 0$  degree and  $2.0^\circ$  and  $2.5^\circ$  off-axis configurations corresponding to the T2K experiment [114]. As expected, an off-axis beam has a narrower spectrum than the 0 degree wide-band beam. Therefore, an off-axis beam is called a “narrow-band beam”. This idea of an off-axis beam was proposed for BNL E889 experiment [113]. It has been employed for the T2K experiment for the first time. Currently, it is also used in the NOvA experiment [115]. For the off-axis beam, obviously the effect of a line neutrino source, namely the difference between the neutrino fluxes measured at the near and far detectors, is enhanced, and it has to be properly taken into account.



**Figure 14.4:** Muon neutrino survival probability at 295 km and neutrino fluxes for different off-axis angles. This figure is taken from Ref. 114.

#### 14.6.4. Reactor neutrino fluxes :

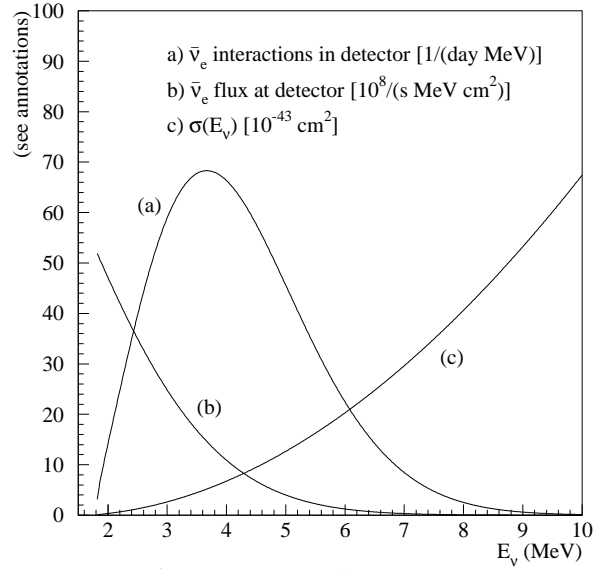
In nuclear reactors, power is generated mainly by nuclear fission of four heavy isotopes,  $^{235}\text{U}$ ,  $^{238}\text{U}$ ,  $^{239}\text{Pu}$ , and  $^{241}\text{Pu}$ . These isotopes account for more than 99% of fissions in the reactor core.  $\beta$ -decays of fission products produce almost pure  $\bar{\nu}_e$  flux. The rate of  $\nu_e$  production is less than  $10^{-5}$  of the rate of  $\bar{\nu}_e$  production [116]. The thermal power outputs of nuclear power reactors are usually quoted in thermal GW,  $\text{GW}_{\text{th}}$ . On the average,  $\sim 200$  MeV and 6 electron antineutrinos are emitted per fission. With 1  $\text{GW}_{\text{th}}$  output,  $\sim 2 \times 10^{20}$  electron antineutrinos are produced per second and emitted isotropically. Typical power plant light-water reactors have thermal power outputs of order 3  $\text{GW}_{\text{th}}$ .

The total  $\bar{\nu}_e$  flux  $S(E_\nu)$  emitted from a reactor is given as a sum of contributions from the four fissioning isotopes,  $S(E_\nu) = \sum_j f_j S_j(E_\nu)$ , where  $f_j$  is the fission rate and  $S_j(E_\nu)$  is the  $\bar{\nu}_e$  flux per fission, of each contributing isotope. The fission rates, and therefore  $S(E_\nu)$ , are

dependent on the thermal power output  $W_{\text{th}}$  from the reactor as a function of time. Using  $W_{\text{th}}$  and the total fission rate  $F = \sum_j f_j$ ,  $S(E_\nu)$  is rewritten as [117]

$$S(E_\nu) = \frac{W_{\text{th}}}{\sum_j (f_j/F) E_j} \sum_j (f_j/F) S_j(E_\nu), \quad (14.25)$$

where  $E_j$  is the energy release per fission by each isotope [118]. The thermal power output and fission fractions  $f_j/F$  are provided by the power plant companies.



**Figure 14.5:** Assuming a 12-ton fiducial mass detector located 0.8 km from 12- $\text{GW}_{\text{th}}$  power reactor,  $\bar{\nu}_e$  interaction spectrum in the detector (curve (a)) and reactor  $\bar{\nu}_e$  flux at the detector (curve (b)) are shown as a function of energy. Inverse  $\beta$ -decay cross section (curve (c)) is also shown. This figure is from Ref. 129.

For  $S_j(E_\nu)$ , a “standard” method used by most of the reactor neutrino oscillation experiments is the one called the “conversion” method [51,119,120], which uses the measured cumulative  $\beta^-$  spectra at ILL for  $^{235}\text{U}$ ,  $^{239}\text{Pu}$ , and  $^{241}\text{Pu}$  [121,122,123]. To estimate the  $\bar{\nu}_e$  spectrum for each isotope, the corresponding  $\beta^-$  spectrum is approximated by the superposition of a set of hypothetical allowed branches. For  $^{238}\text{U}$ , theoretical calculations [124] have been used. Recently, however, the cumulative  $\beta^-$  spectrum for  $^{238}\text{U}$  has been measured at the scientific neutron source FRM II in Garching, Germany, and the converted  $\bar{\nu}_e$  spectrum is given for  $2.875 < E_\nu < 7.625$  MeV [125]. It is used in the recent neutrino oscillation analysis of the Double Chooz experiment [40].

While the conversion of the  $\beta$  spectrum to the  $\bar{\nu}_e$  spectrum is trivial for a single  $\beta$ -decay branch, fission of the four main fuel isotopes involve  $> 1000$  daughter isotopes and  $> 6000$  individual  $\beta$ -decay branches (approximately six thousands), thus causing rather large uncertainties in both the normalization and shape of the reactor  $\bar{\nu}_e$  flux. Recent detailed reactor  $\bar{\nu}_e$  calculations [119,120] show a few % larger fluxes with respect to the fluxes calculated in Ref. 122. The antineutrino fluxes measured by previous short baseline reactor neutrino oscillation experiments are generally in agreement with the latter calculation. The recent reactor antineutrino flux measurement at Daya Bay [126] is also consistent with these previous measurements. This discrepancy is called the “reactor neutrino anomaly” and hints to possible oscillations involving sterile neutrinos (see Section 14.13). For the shape of the reactor  $\bar{\nu}_e$  flux, all the current reactor neutrino oscillation experiments, Daya Bay [126], RENO [39,127], and Double Chooz [40], observe an excess of  $\bar{\nu}_e$  flux in the energy region from 4 to 6 MeV, relative to current predictions. The excess rate is observed to

be time independent and correlated with the reactor power. Because of this, an unknown background is unlikely for its explanation. There are certain suggestions on the possible origins of this excess, but this problem is not completely solved yet [128].

Electron antineutrinos from reactors are detected via the inverse  $\beta$ -decay  $\bar{\nu}_e + p \rightarrow e^+ + n$ . This reaction has a threshold of 1.8 MeV, so that the  $\bar{\nu}_e$  flux above this threshold is detected. The event rate as a function of  $\bar{\nu}_e$  energy  $E_\nu$  is proportional to  $\sigma(E_\nu)S(E_\nu)$ , where  $\sigma(E_\nu)$  is the cross section of the inverse  $\beta$ -decay. Fig. 14.5 shows  $\sigma(E_\nu)$ . This figure also shows the flux and event rate for a particular detector configuration (see caption to this figure) in a reactor neutrino oscillation experiment.

## 14.7. Neutrino oscillations in vacuum

Neutrino oscillations are a quantum mechanical consequence of the existence of nonzero neutrino masses and neutrino (lepton) mixing, Eq. (14.1), and of the relatively small splitting between the neutrino masses. The neutrino mixing and oscillation phenomena are analogous to the  $K^0 - \bar{K}^0$  and  $B^0 - \bar{B}^0$  mixing and oscillations.

In what follows we will present a simplified version of the derivation of the expressions for the neutrino and antineutrino oscillation probabilities. The complete derivation would require the use of the wave packet formalism for the evolution of the massive neutrino states, or, alternatively, of the field-theoretical approach, in which one takes into account the processes of production, propagation and detection of neutrinos [130].

Suppose the flavour neutrino  $\nu_l$  is produced in a CC weak interaction process and after a time  $T$  it is observed by a neutrino detector, located at a distance  $L$  from the neutrino source and capable of detecting also neutrinos  $\nu_{l'}$ ,  $l' \neq l$ . We will consider the evolution of the neutrino state  $|\nu_l\rangle$  in the frame in which the detector is at rest (laboratory frame). The oscillation probability, as we will see, is a Lorentz invariant quantity. If lepton mixing, Eq. (14.1), takes place and the masses  $m_j$  of all neutrinos  $\nu_j$  are sufficiently small, the state of the neutrino  $\nu_l$ ,  $|\nu_l\rangle$ , will be a coherent superposition of the states  $|\nu_j\rangle$  of neutrinos  $\nu_j$ :

$$|\nu_l\rangle = \sum_j U_{lj}^* |\nu_j; \tilde{p}_j\rangle, \quad l = e, \mu, \tau, \quad (14.26)$$

where  $U$  is the neutrino mixing matrix and  $\tilde{p}_j$  is the 4-momentum of  $\nu_j$  [131].

We will consider the case of relativistic neutrinos  $\nu_j$ , which corresponds to the conditions in both past and currently planned future neutrino oscillation experiments [133]. In this case the state  $|\nu_j; \tilde{p}_j\rangle$  practically coincides with the helicity (-1) state  $|\nu_j, L; \tilde{p}_j\rangle$  of the neutrino  $\nu_j$ , the admixture of the helicity (+1) state  $|\nu_j, R; \tilde{p}_j\rangle$  in  $|\nu_j; \tilde{p}_j\rangle$  being suppressed due to the factor  $\sim m_j/E_j$ , where  $E_j$  is the energy of  $\nu_j$ . If  $\nu_j$  are Majorana particles,  $\nu_j \equiv \chi_j$ , due to the presence of the helicity (+1) state  $|\chi_j, R; \tilde{p}_j\rangle$  in  $|\chi_j; \tilde{p}_j\rangle$ , the neutrino  $\nu_l$  can produce an  $l^+$  (instead of  $l^-$ ) when it interacts, *e.g.*, with nucleons. The cross section of such a  $|\Delta L_l| = 2$  process is suppressed by the factor  $(m_j/E_j)^2$ , which renders the process unobservable at present.

If the number  $n$  of massive neutrinos  $\nu_j$  is bigger than 3 due to a mixing between the active flavour and sterile neutrinos, one will have additional relations similar to that in Eq. (14.26) for the state vectors of the (predominantly LH) sterile antineutrinos. In the case of just one RH sterile neutrino field  $\nu_{sR}(x)$ , for instance, we will have in addition to Eq. (14.26):

$$|\bar{\nu}_{sL}\rangle = \sum_{j=1}^4 U_{sj}^* |\nu_j; \tilde{p}_j\rangle \cong \sum_{j=1}^4 U_{sj}^* |\nu_j, L; \tilde{p}_j\rangle, \quad (14.27)$$

where the neutrino mixing matrix  $U$  is now a  $4 \times 4$  unitary matrix.

For the state vector of RH flavour antineutrino  $\bar{\nu}_l$ , produced in a CC weak interaction process we similarly get:

$$|\bar{\nu}_l\rangle = \sum_j U_{lj} |\bar{\nu}_j; \tilde{p}_j\rangle \cong \sum_{j=1}^4 U_{lj} |\bar{\nu}_j, R; \tilde{p}_j\rangle, \quad l = e, \mu, \tau, \quad (14.28)$$

where  $|\bar{\nu}_j, R; \tilde{p}_j\rangle$  is the helicity (+1) state of the antineutrino  $\bar{\nu}_j$  if  $\nu_j$  are Dirac fermions, or the helicity (+1) state of the neutrino  $\nu_j \equiv \bar{\nu}_j \equiv \chi_j$  if the massive neutrinos are Majorana particles. Thus, in the latter case we have in Eq. (14.28):  $|\bar{\nu}_j; \tilde{p}_j\rangle \cong |\nu_j, R; \tilde{p}_j\rangle \equiv |\chi_j, R; \tilde{p}_j\rangle$ . The presence of the matrix  $U$  in Eq. (14.28) (and not of  $U^*$ ) follows directly from Eq. (14.1).

We will assume in what follows that the spectrum of masses of neutrinos is not degenerate:  $m_j \neq m_k$ ,  $j \neq k$ . Then the states  $|\nu_j; \tilde{p}_j\rangle$  in the linear superposition in the r.h.s. of Eq. (14.26) will have, in general, different energies and different momenta, independently of whether they are produced in a decay or interaction process:  $\tilde{p}_j \neq \tilde{p}_k$ , or  $E_j \neq E_k$ ,  $\mathbf{p}_j \neq \mathbf{p}_k$ ,  $j \neq k$ , where  $E_j = \sqrt{p_j^2 + m_j^2}$ ,  $p_j \equiv |\mathbf{p}_j|$ . The deviations of  $E_j$  and  $p_j$  from the values for a massless neutrino  $E$  and  $p = E$  are proportional to  $m_j^2/E_0$ ,  $E_0$  being a characteristic energy of the process, and are extremely small. In the case of  $\pi^+ \rightarrow \mu^+ + \nu_\mu$  decay at rest, for instance, we have:  $E_j = E + m_j^2/(2m_\pi)$ ,  $p_j = E - \xi m_j^2/(2E)$ , where  $E = (m_\pi/2)(1 - m_\mu^2/m_\pi^2) \cong 30$  MeV,  $\xi = (1 + m_\mu^2/m_\pi^2)/2 \cong 0.8$ , and  $m_\mu$  and  $m_\pi$  are the  $\mu^+$  and  $\pi^+$  masses. Taking  $m_j = 1$  eV we find:  $E_j \cong E(1 + 1.2 \times 10^{-16})$  and  $p_j \cong E(1 - 4.4 \times 10^{-16})$ .

Given the uncorrelated uncertainties  $\delta E$  and  $\delta p$  in the knowledge of the neutrino energy  $E$  and momentum  $p$ , the quantum mechanical condition that neutrinos with definite mass  $\nu_1, \nu_2, \dots$ , whose states are part of the linear superposition of states corresponding, for example, to  $|\nu_l\rangle$  in Eq. (14.26), are emitted coherently when  $|\nu_l\rangle$  is produced in some weak interaction process, has the form [134]:

$$\delta m^2 = \sqrt{(2E\delta E)^2 + (2p\delta p)^2} > \max(|m_i^2 - m_j^2|), \quad i, j = 1, 2, \dots, n, \quad (14.29)$$

where  $\delta m^2$  is the uncertainty in the square of the neutrino mass due to the uncertainties in the energy and momentum of the neutrino. Equation Eq. (14.29) follows from the well known relativistic relation  $E^2 = p^2 + m^2$ . In the context under discussion,  $\delta E$  and  $\delta p$  should be understood as the intrinsic quantum mechanical uncertainties in the neutrino energy and momentum for the given neutrino production and detection processes, *i.e.*,  $\delta E$  and  $\delta p$  are the minimal uncertainties with which  $E$  and  $p$  can be determined in the considered production and detection processes. Then  $\delta m^2$  is the quantum mechanical uncertainty of the inferred squared neutrino mass.

Suppose that the neutrinos are observed via a CC weak interaction process and that in the detector's rest frame they are detected after time  $T$  after emission, after traveling a distance  $L$ . Then the amplitude of the probability that neutrino  $\nu_{l'}$  will be observed if neutrino  $\nu_l$  was produced by the neutrino source can be written as [130,132,135]:

$$A(\nu_l \rightarrow \nu_{l'}) = \sum_j U_{l'j} D_j U_{lj}^\dagger, \quad l, l' = e, \mu, \tau, \quad (14.30)$$

where  $D_j = D_j(p_j; L, T)$  describes the propagation of  $\nu_j$  between the source and the detector,  $U_{jl}^\dagger$  and  $U_{l'j}$  are the amplitudes to find  $\nu_j$  in the initial and in the final flavour neutrino state, respectively. It follows from relativistic Quantum Mechanics considerations that [130,132]

$$D_j \equiv D_j(\tilde{p}_j; L, T) = e^{-i\tilde{p}_j(x_f - x_0)} = e^{-i(E_j T - p_j L)}, \quad p_j \equiv |\mathbf{p}_j|, \quad (14.31)$$

where [136]  $x_0$  and  $x_f$  are the space-time coordinates of the points of neutrino production and detection,  $T = (t_f - t_0)$  and  $L = \mathbf{k}(x_f - x_0)$ ,  $\mathbf{k}$  being the unit vector in the direction of neutrino momentum,  $\mathbf{p}_j = k\mathbf{p}_j$ . What is relevant for the calculation of the probability  $P(\nu_l \rightarrow \nu_{l'}) = |A(\nu_l \rightarrow \nu_{l'})|^2$  is the interference factor  $D_j D_k^*$  which depends on the phase

$$\begin{aligned} \delta\varphi_{jk} &= (E_j - E_k)T - (p_j - p_k)L = (E_j - E_k) \left[ T - \frac{E_j + E_k}{p_j + p_k} L \right] \\ &+ \frac{m_j^2 - m_k^2}{p_j + p_k} L. \end{aligned} \quad (14.32)$$

Some authors [137] have suggested that the distance traveled by the neutrinos  $L$  and the time interval  $T$  are related by  $T = (E_j + E_k) L / (p_j + p_k) = L / \bar{v}$ ,  $\bar{v} = (E_j / (E_j + E_k)) v_j + (E_k / (E_j + E_k)) v_k$  being the “average” velocity of  $\nu_j$  and  $\nu_k$ , where  $v_{j,k} = p_{j,k} / E_{j,k}$ . In this case the first term in the r.h.s. of Eq. (14.32) vanishes. The indicated relation has not emerged so far from any dynamical wave packet calculations. We arrive at the same conclusion concerning the term under discussion in Eq. (14.32) if one assumes [138] that  $E_j = E_k = E_0$ . Finally, it was proposed in Ref. 135 and Ref. 139 that the states of  $\nu_j$  and  $\bar{\nu}_j$  in Eq. (14.26) and Eq. (14.28) have the same 3-momentum,  $p_j = p_k = p$ . Under this condition the first term in the r.h.s. of Eq. (14.32) is negligible, being suppressed by the additional factor  $(m_j^2 + m_k^2) / p^2$  since for relativistic neutrinos  $L = T$  up to terms  $\sim m_{j,k}^2 / p^2$ . We arrive at the same conclusion if  $E_j \neq E_k$ ,  $p_j \neq p_k$ ,  $j \neq k$ , and we take into account that neutrinos are relativistic and therefore, up to corrections  $\sim m_{j,k}^2 / E_{j,k}^2$ , we have  $L \cong T$  (see, *e.g.*, C. Giunti quoted in Ref. 130).

Although the cases considered above are physically quite different, they lead to the same result for the phase difference  $\delta\varphi_{jk}$ . Thus, we have:

$$\delta\varphi_{jk} \cong \frac{m_j^2 - m_k^2}{2p} L = 2\pi \frac{L}{L_{jk}^v} \text{sgn}(m_j^2 - m_k^2), \quad (14.33)$$

where  $p = (p_j + p_k) / 2$  and

$$L_{jk}^v = 4\pi \frac{p}{|\Delta m_{jk}^2|} \cong 2.48 \text{ m} \frac{p[\text{MeV}]}{|\Delta m_{jk}^2|[\text{eV}^2]} \quad (14.34)$$

is the neutrino oscillation length associated with  $\Delta m_{jk}^2$ . We can safely neglect the dependence of  $p_j$  and  $p_k$  on the masses  $m_j$  and  $m_k$  and consider  $p$  to be the zero neutrino mass momentum,  $p = E$ . The phase difference  $\delta\varphi_{jk}$ , Eq. (14.33), is Lorentz-invariant.

Eq. (14.31) corresponds to a plane-wave description of the propagation of neutrinos  $\nu_j$ . It accounts only for the movement of the center of the wave packet describing  $\nu_j$ . In the wave packet treatment of the problem, the interference between the states of  $\nu_j$  and  $\nu_k$  is subject to a number of conditions [130], the localization condition and the condition of overlapping of the wave packets of  $\nu_j$  and  $\nu_k$  at the detection point being the most important. For relativistic neutrinos, the localisation condition in space, for instance, reads:  $\sigma_{xP}, \sigma_{xD} < L_{jk}^v / (2\pi)$ ,  $\sigma_{xP(D)}$  being the spatial width of the production (detection) wave packet. Thus, the interference will not be suppressed if the spatial width of the neutrino wave packets determined by the neutrino production and detection processes is smaller than the corresponding oscillation length in vacuum. In order for the interference to be nonzero, the wave packets describing  $\nu_j$  and  $\nu_k$  should also overlap in the point of neutrino detection. This requires that the spatial separation between the two wave packets at the point of neutrinos detection, caused by the two wave packets having different group velocities  $v_j \neq v_k$ , satisfies  $|(v_j - v_k)T| \ll \max(\sigma_{xP}, \sigma_{xD})$ . If the interval of time  $T$  is not measured,  $T$  in the preceding condition must be replaced by the distance  $L$  between the neutrino source and the detector (for further discussion see, *e.g.*, Refs. [130,132,135]).

For the  $\nu_l \rightarrow \nu_{l'}$  and  $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$  oscillation probabilities we get from Eq. (14.30), Eq. (14.31), and Eq. (14.33):

$$P(\nu_l \rightarrow \nu_{l'}) = \sum_j |U_{lj}|^2 |U_{l'j}|^2 + 2 \sum_{j>k} |U_{lj} U_{l'j}^* U_{lk} U_{l'k}^*| \cos\left(\frac{\Delta m_{jk}^2}{2p} L - \phi_{l'l';jk}\right), \quad (14.35)$$

$$P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'}) = \sum_j |U_{lj}|^2 |U_{l'j}|^2 + 2 \sum_{j>k} |U_{lj} U_{l'j}^* U_{lk} U_{l'k}^*| \cos\left(\frac{\Delta m_{jk}^2}{2p} L + \phi_{l'l';jk}\right), \quad (14.36)$$

where  $l, l' = e, \mu, \tau$  and  $\phi_{l'l';jk} = \arg(U_{lj} U_{l'j}^* U_{lk} U_{l'k}^*)$ . It follows from Eq. (14.30) - Eq. (14.32) that in order for neutrino oscillations

to occur, at least two neutrinos  $\nu_j$  should not be degenerate in mass and lepton mixing should take place,  $U \neq \mathbf{1}$ . The neutrino oscillations effects can be large if we have

$$\frac{|\Delta m_{jk}^2|}{2p} L = 2\pi \frac{L}{L_{jk}^v} \gtrsim 1, \quad j \neq k. \quad (14.37)$$

at least for one  $\Delta m_{jk}^2$ . This condition has a simple physical interpretation: the neutrino oscillation length  $L_{jk}^v$  should be of the order of, or smaller, than source-detector distance  $L$ , otherwise the oscillations will not have time to develop before neutrinos reach the detector.

We see from Eq. (14.35) and Eq. (14.36) that  $P(\nu_l \rightarrow \nu_{l'}) = P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$ ,  $l, l' = e, \mu, \tau$ . This is a consequence of CPT invariance. The conditions of CP and T invariance read [48,61,62]:  $P(\nu_l \rightarrow \nu_{l'}) = P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$ ,  $l, l' = e, \mu, \tau$  (CP),  $P(\nu_l \rightarrow \nu_{l'}) = P(\nu_{l'} \rightarrow \nu_l)$ ,  $P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'}) = P(\bar{\nu}_{l'} \rightarrow \bar{\nu}_l)$ ,  $l, l' = e, \mu, \tau$  (T). In the case of CPT invariance, which we will assume to hold throughout this article, we get for the survival probabilities:  $P(\nu_l \rightarrow \nu_l) = P(\bar{\nu}_l \rightarrow \bar{\nu}_l)$ ,  $l, l' = e, \mu, \tau$ . Thus, the study of the “disappearance” of  $\nu_l$  and  $\bar{\nu}_l$ , caused by oscillations in vacuum, cannot be used to test whether CP invariance holds in the lepton sector. It follows from Eq. (14.35) and Eq. (14.36) that we can have CP violation effects in neutrino oscillations only if  $\phi_{l'l';jk} \neq \pi q$ ,  $q = 0, 1, 2$ , *i.e.*, if  $U_{lj} U_{l'j}^* U_{lk} U_{l'k}^* \neq 1$ , and therefore  $U$  itself, is not real. As a measure of CP and T violation in neutrino oscillations we can consider the asymmetries:

$$A_{\text{CP}}^{(l'l)} \equiv P(\nu_l \rightarrow \nu_{l'}) - P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'}), \quad A_{\text{T}}^{(l'l)} \equiv P(\nu_l \rightarrow \nu_{l'}) - P(\nu_{l'} \rightarrow \nu_l). \quad (14.38)$$

CPT invariance implies:  $A_{\text{CP}}^{(l'l)} = -A_{\text{CP}}^{(l'l)}$ ,  $A_{\text{T}}^{(l'l)} = P(\bar{\nu}_{l'} \rightarrow \bar{\nu}_l) - P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'}) = A_{\text{CP}}^{(l'l)}$ . It follows further directly from Eq. (14.35) and Eq. (14.36) that

$$A_{\text{CP}}^{(l'l)} = 4 \sum_{j>k} \text{Im} \left( U_{lj} U_{l'j}^* U_{lk} U_{l'k}^* \right) \sin \frac{\Delta m_{jk}^2}{2p} L, \quad l, l' = e, \mu, \tau. \quad (14.39)$$

Eq. (14.2) and Eq. (14.35) - Eq. (14.36) imply that  $P(\nu_l \rightarrow \nu_{l'})$  and  $P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$  do not depend on the Majorana CP violation phases in the neutrino mixing matrix  $U$  [48]. Thus, the experiments investigating the  $\nu_l \rightarrow \nu_{l'}$  and  $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$  oscillations,  $l, l' = e, \mu, \tau$ , cannot provide information on the nature - Dirac or Majorana, of massive neutrinos. The same conclusions hold also when the  $\nu_l \rightarrow \nu_{l'}$  and  $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$  oscillations take place in matter [65]. In the case of  $\nu_l \leftrightarrow \nu_{l'}$  and  $\bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$  oscillations in vacuum, only the Dirac phase(s) in  $U$  can cause CP violating effects leading to  $P(\nu_l \rightarrow \nu_{l'}) \neq P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$ ,  $l \neq l'$ .

In the case of 3-neutrino mixing all different  $\text{Im}(U_{lj} U_{l'j}^* U_{lk} U_{l'k}^*) \neq 0$ ,  $l' \neq l = e, \mu, \tau$ ,  $j \neq k = 1, 2, 3$ , coincide up to a sign as a consequence of the unitarity of  $U$ . Therefore one has [63]:

$$A_{\text{CP}}^{(\mu e)} = -A_{\text{CP}}^{(\tau e)} = A_{\text{CP}}^{(\tau \mu)} = 4 J_{\text{CP}} \left( \sin \frac{\Delta m_{32}^2}{2p} L + \sin \frac{\Delta m_{21}^2}{2p} L + \sin \frac{\Delta m_{13}^2}{2p} L \right) \quad (14.40)$$

where

$$J_{\text{CP}} = \text{Im} \left( U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^* \right), \quad (14.41)$$

is the “rephasing invariant” associated with the Dirac CP violation phase in  $U$ . It is analogous to the rephasing invariant associated with the Dirac CP violating phase in the CKM quark mixing matrix [64]. It is clear from Eq. (14.40) that  $J_{\text{CP}}$  controls the magnitude of CP violation effects in neutrino oscillations in the case of 3-neutrino mixing. If  $\sin(\Delta m_{ij}^2 / (2p)L) \cong 0$  for  $(ij) = (32)$ , or  $(21)$ , or  $(13)$ ,

we get  $A_{\text{CP}}^{(l'l)} \cong 0$ . Thus, if as a consequence of the production, propagation and/or detection of neutrinos, effectively oscillations due only to one non-zero neutrino mass squared difference take place, the CP violating effects will be strongly suppressed. In particular, we get  $A_{\text{CP}}^{(l'l)} = 0$ , unless all three  $\Delta m_{ij}^2 \neq 0$ ,  $(ij) = (32), (21), (13)$ .



If the number of massive neutrinos  $n$  is equal to the number of neutrino flavours,  $n = 3$ , one has as a consequence of the unitarity of the neutrino mixing matrix:  $\sum_{l'=e,\mu,\tau} P(\nu_l \rightarrow \nu_{l'}) = 1$ ,  $l = e, \mu, \tau$ ,  $\sum_{l=e,\mu,\tau} P(\nu_l \rightarrow \nu_{l'}) = 1$ ,  $l' = e, \mu, \tau$ . Similar “probability conservation” equations hold for  $P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$ . If, however, the number of light massive neutrinos is bigger than the number of flavour neutrinos as a consequence, *e.g.*, of a flavour neutrino - sterile neutrino mixing, we would have  $\sum_{l'=e,\mu,\tau} P(\nu_l \rightarrow \nu_{l'}) = 1 - P(\nu_l \rightarrow \bar{\nu}_{sL})$ ,  $l = e, \mu, \tau$ , where we have assumed the existence of just one sterile neutrino. Obviously, in this case  $\sum_{l'=e,\mu,\tau} P(\nu_l \rightarrow \nu_{l'}) < 1$  if  $P(\nu_l \rightarrow \bar{\nu}_{sL}) \neq 0$ . The former inequality is used in the searches for oscillations between active and sterile neutrinos.

Consider next neutrino oscillations in the case of one neutrino mass squared difference “dominance”: suppose that  $|\Delta m_{j1}^2| \ll |\Delta m_{n1}^2|$ ,  $j = 2, \dots, (n-1)$ ,  $|\Delta m_{n1}^2| L/(2p) \gtrsim 1$  and  $|\Delta m_{j1}^2| L/(2p) \ll 1$ , so that  $\exp[i(\Delta m_{j1}^2 L/(2p))] \cong 1$ ,  $j = 2, \dots, (n-1)$ . Under these conditions we obtain from Eq. (14.35) and Eq. (14.36), keeping only the oscillating terms involving  $\Delta m_{n1}^2$ :

$$P(\nu_{l(l')} \rightarrow \nu_{l'(l)}) \cong P(\bar{\nu}_{l(l)} \rightarrow \bar{\nu}_{l'(l)}) \cong \delta_{ll'} - 2|U_{ln}|^2 \left[ \delta_{ll'} - |U_{l'n}|^2 \right] (1 - \cos \frac{\Delta m_{n1}^2}{2p} L). \quad (14.42)$$

It follows from the neutrino oscillation data discussed in Section 14.2 that in the case of 3-neutrino mixing, one of the two independent neutrino mass squared differences,  $\Delta m_{21}^2 > 0$ , is much smaller than the absolute value of the second one,  $|\Delta m_{31}^2|$ :  $\Delta m_{21}^2 \ll |\Delta m_{31}^2|$ . The data, as we have seen, imply:

$$\begin{aligned} \Delta m_{21}^2 &\cong 7.4 \times 10^{-5} \text{ eV}^2, \\ |\Delta m_{31}^2| &\cong 2.5 \times 10^{-3} \text{ eV}^2, \\ \Delta m_{21}^2/|\Delta m_{31}^2| &\cong 0.03. \end{aligned} \quad (14.43)$$

Neglecting the effects due to  $\Delta m_{21}^2$  we get from Eq. (14.42) by setting  $n = 3$  and choosing, *e.g.*, i)  $l = l' = e$  and ii)  $l = e(\mu)$ ,  $l' = \mu(e)$  [140]:

$$P(\nu_e \rightarrow \nu_e) = P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \cong 1 - 2|U_{e3}|^2 (1 - |U_{e3}|^2) \left( 1 - \cos \frac{\Delta m_{31}^2}{2p} L \right), \quad (14.44)$$

$$\begin{aligned} P(\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}) &\cong 2|U_{\mu3}|^2 |U_{e3}|^2 \left( 1 - \cos \frac{\Delta m_{31}^2}{2p} L \right) \\ &= \frac{|U_{\mu3}|^2}{1 - |U_{e3}|^2} P^{2\nu}(|U_{e3}|^2, m_{31}^2), \end{aligned} \quad (14.45)$$

and  $P(\bar{\nu}_{\mu(e)} \rightarrow \bar{\nu}_{e(\mu)}) = P(\nu_{\mu(e)} \rightarrow \nu_{e(\mu)})$ . Here  $P^{2\nu}(|U_{e3}|^2, m_{31}^2)$  is the probability of the 2-neutrino transition  $\nu_e \rightarrow (\nu_{23} + \nu_{32})$  due to  $\Delta m_{31}^2$  and a mixing with angle  $\theta_{13}$ , where

$$\begin{aligned} \sin^2 \theta_{13} &= |U_{e3}|^2, \quad s_{23}^2 \equiv \sin^2 \theta_{23} = \frac{|U_{\mu3}|^2}{1 - |U_{e3}|^2}, \\ c_{23}^2 &\equiv \cos^2 \theta_{23} = \frac{|U_{\tau3}|^2}{1 - |U_{e3}|^2}, \end{aligned} \quad (14.46)$$

*i.e.*,  $\theta_{13}$  and  $\theta_{23}$  are the angles of the standard parametrization of the neutrino mixing matrix. Eq. (14.44) describes with a relatively high precision the oscillations of reactor  $\bar{\nu}_e$  on a distance  $L \sim 1$  km in the case of 3-neutrino mixing. It was used in the analysis of the data of the Chooz [51], Double Chooz [30], Daya Bay [31] and RENO [32] experiments. Eq. (14.42) with  $n = 3$  and  $l = l' = \mu$  describes with a relatively good precision the effects of “disappearance” due to oscillations of the accelerator  $\nu_\mu$ , seen in the K2K [19] MINOS [20,21] and T2K [22,23] experiments. The  $\nu_\mu \rightarrow \nu_\tau$  transitions due to the oscillations, which the OPERA experiment [141,142] is observing, can be described by Eq. (14.42) with  $n = 3$  and  $l = \mu$ ,  $l' = \tau$ . Finally, the probability Eq. (14.45) describes with a good precision the  $\nu_\mu \rightarrow \nu_e$  and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations under the conditions of the K2K experiment.

In certain cases the dimensions of the neutrino source,  $\Delta L$ , are not negligible in comparison with the oscillation length. Similarly,

when analyzing neutrino oscillation data one has to include the energy resolution of the detector,  $\Delta E$ , *etc.* in the analysis. As can be shown [46], if  $2\pi\Delta L/L_{jk}^v \gg 1$ , and/or  $2\pi(L/L_{jk}^v)(\Delta E/E) \gg 1$ , the oscillating terms in the neutrino oscillation probabilities will be strongly suppressed. In this case (as well as in the case of sufficiently large separation of the  $\nu_j$  and  $\nu_k$  wave packets at the detection point) the interference terms in  $P(\nu_l \rightarrow \nu_{l'})$  and  $P(\bar{\nu}_{l'} \rightarrow \bar{\nu}_l)$  will be negligibly small and the neutrino flavour conversion will be determined by the average probabilities:

$$\bar{P}(\nu_l \rightarrow \nu_{l'}) = \bar{P}(\bar{\nu}_l \rightarrow \bar{\nu}_{l'}) \cong \sum_j |U_{lj}|^2 |U_{lj}|^2. \quad (14.47)$$

Suppose next that in the case of 3-neutrino mixing,  $|\Delta m_{21}^2| L/(2p) \sim 1$ , while at the same time  $|\Delta m_{31(32)}^2| L/(2p) \gg 1$ , and the oscillations due to  $\Delta m_{31}^2$  and  $\Delta m_{32}^2$  are strongly suppressed (averaged out) due to integration over the region of neutrino production, the energy resolution function, *etc.* In this case we get for the  $\nu_e$  and  $\bar{\nu}_e$  survival probabilities:

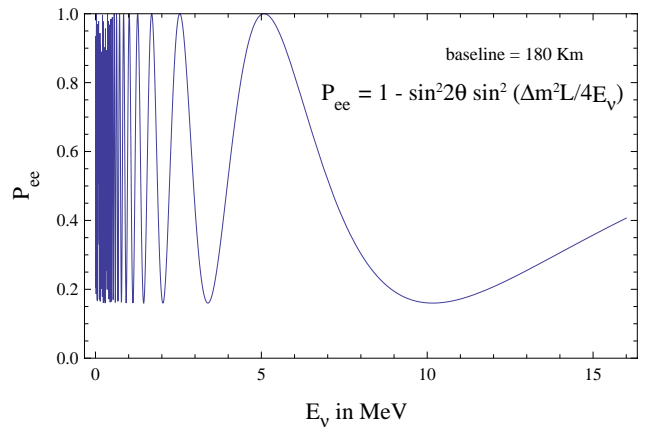
$$P(\nu_e \rightarrow \nu_e) = P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \cong |U_{e3}|^4 + (1 - |U_{e3}|^2)^2 P^{2\nu}(\nu_e \rightarrow \nu_e), \quad (14.48)$$

$$\begin{aligned} P^{2\nu}(\nu_e \rightarrow \nu_e) &= P^{2\nu}(\bar{\nu}_e \rightarrow \bar{\nu}_e) \equiv P_{ee}^{2\nu}(\theta_{12}, \Delta m_{21}^2) \\ &= 1 - \frac{1}{2} \sin^2 2\theta_{12} \left( 1 - \cos \frac{\Delta m_{21}^2}{2p} L \right) \\ &= 1 - \sin^2 2\theta_{12} \sin^2 \left( \frac{\Delta m_{21}^2}{4E} L \right) \end{aligned} \quad (14.49)$$

being the  $\nu_e$  and  $\bar{\nu}_e$  survival probability in the case of 2-neutrino oscillations “driven” by the angle  $\theta_{12}$  and  $\Delta m_{21}^2$ , with  $\theta_{12}$  given by

$$\cos^2 \theta_{12} = \frac{|U_{e1}|^2}{1 - |U_{e3}|^2}, \quad \sin^2 \theta_{12} = \frac{|U_{e2}|^2}{1 - |U_{e3}|^2}. \quad (14.50)$$

Eq. (14.48) with  $P^{2\nu}(\bar{\nu}_e \rightarrow \bar{\nu}_e)$  given by Eq. (14.49) describes the effects of neutrino oscillations of reactor  $\bar{\nu}_e$  observed by the KamLAND experiment.



**Figure 14.6:** The  $\nu_e$  ( $\bar{\nu}_e$ ) survival probability  $P(\nu_e \rightarrow \nu_e) = P(\bar{\nu}_e \rightarrow \bar{\nu}_e)$ , Eq. (14.52), as a function of the neutrino energy for  $L = 180$  km,  $\Delta m^2 = 7.0 \times 10^{-5} \text{ eV}^2$  and  $\sin^2 2\theta = 0.84$  (from Ref. 143).

The data of  $\nu$ -oscillations experiments were often analyzed in the past, and in certain cases new data are still analyzed at present, assuming 2-neutrino mixing:

$$|\nu_l\rangle = |\nu_1\rangle \cos \theta + |\nu_2\rangle \sin \theta, \quad |\nu_x\rangle = -|\nu_1\rangle \sin \theta + |\nu_2\rangle \cos \theta, \quad (14.51)$$

where  $\theta$  is the neutrino mixing angle in vacuum and  $\nu_x$  is another flavour neutrino or sterile (anti-) neutrino,  $x = l' \neq l$  or  $\nu_x \equiv \bar{\nu}_{sL}$ . In this case we have [139]:

$$\begin{aligned} P^{2\nu}(\nu_l \rightarrow \nu_l) &= 1 - \frac{1}{2} \sin^2 2\theta \left( 1 - \cos 2\pi \frac{L}{L^v} \right) \\ &= 1 - \sin^2 2\theta \left( \sin^2 \frac{\Delta m^2}{4E} L \right), \\ P^{2\nu}(\nu_l \rightarrow \nu_x) &= 1 - P^{2\nu}(\nu_l \rightarrow \nu_l), \end{aligned} \quad (14.52)$$

where  $L^v = 4\pi E / \Delta m^2$  ( $p = E$ ),  $\Delta m^2 = m_2^2 - m_1^2 > 0$ . Combining the CPT invariance constraints with the probability conservation one obtains:  $P(\nu_l \rightarrow \nu_x) = P(\bar{\nu}_l \rightarrow \bar{\nu}_x) = P(\nu_x \rightarrow \nu_l) = P(\bar{\nu}_x \rightarrow \bar{\nu}_l)$ . These equalities and Eq. (14.52) with  $l = \mu$  and  $x = \tau$  were used, for instance, in the analysis of the Super-K atmospheric neutrino data [17], in which the first compelling evidence for oscillations of neutrinos was obtained. The probability  $P^{2\nu}(\nu_l \rightarrow \nu_x)$ , Eq. (14.52), depends on two factors: on  $(1 - \cos 2\pi L / L^v)$ , which exhibits oscillatory dependence on the distance  $L$  and on the neutrino energy  $p = E$  (hence the name “neutrino oscillations”), and on  $\sin^2 2\theta$ , which determines the amplitude of the oscillations. In order to have  $P^{2\nu}(\nu_l \rightarrow \nu_x) \cong 1$ , two conditions have to be fulfilled: one should have  $\sin^2 2\theta \cong 1$  and  $L^v \lesssim 2\pi L$  with  $\cos 2\pi L / L^v \cong -1$ . If  $L^v \gg 2\pi L$ , the oscillations do not have enough time to develop on the way to the neutrino detector and  $P(\nu_l \rightarrow \nu_x) \cong 0$ , while  $P(\nu_l \rightarrow \nu_l) \cong 1$ . The preceding comments are illustrated in Fig. 14.6 showing the dependence of the probability  $P^{2\nu}(\nu_e \rightarrow \nu_e) = P^{2\nu}(\bar{\nu}_e \rightarrow \bar{\nu}_e)$  on the neutrino energy.

**Table 14.4:** Sensitivity of different oscillation experiments.

Source	Type of $\nu$	$\bar{E}$ [MeV]	$L$ [km]	$\min(\Delta m^2)$ [eV <sup>2</sup> ]
Reactor	$\bar{\nu}_e$	$\sim 1$	1	$\sim 10^{-3}$
Reactor	$\bar{\nu}_e$	$\sim 1$	100	$\sim 10^{-5}$
Accelerator	$\nu_\mu, \bar{\nu}_\mu$	$\sim 10^3$	1	$\sim 1$
Accelerator	$\nu_\mu, \bar{\nu}_\mu$	$\sim 10^3$	1000	$\sim 10^{-3}$
Atmospheric $\nu$ 's	$\nu_{\mu,e}, \bar{\nu}_{\mu,e}$	$\sim 10^3$	$10^4$	$\sim 10^{-4}$
Sun	$\nu_e$	$\sim 1$	$1.5 \times 10^8$	$\sim 10^{-11}$

A given experiment searching for neutrino oscillations is specified, in particular, by the average energy of the neutrinos being studied,  $\bar{E}$ , and by the source-detector distance  $L$ . The requirement  $L_{jk}^v \lesssim 2\pi L$  determines the minimal value of a generic neutrino mass squared difference  $\Delta m^2 > 0$ , to which the experiment is sensitive (figure of merit of the experiment):  $\min(\Delta m^2) \sim 2\bar{E}/L$ . Because of the interference nature of neutrino oscillations, experiments can probe, in general, rather small values of  $\Delta m^2$  (see, e.g., Ref. 135). Values of  $\min(\Delta m^2)$ , characterizing qualitatively the sensitivity of different experiments are given in Table 14.4. They correspond to the reactor experiments Chooz, Daya Bay, RENO, Double Chooz ( $L \sim 1$  km) and KamLAND ( $L \sim 100$  km), to accelerator experiments - past ( $L \sim 1$  km), and current (K2K, MINOS, OPERA, T2K, NO $\nu$ A [115]),  $L \sim (300 \div 1000)$  km), to the Super-Kamiokande, MINOS and IceCube-DeepCore experiments studying atmospheric neutrino oscillations, and to the solar neutrino experiments.

#### 14.8. Matter effects in neutrino oscillations

The presence of matter can change drastically the pattern of neutrino oscillations: neutrinos can interact with the particles forming the matter. Accordingly, the Hamiltonian of the neutrino system in matter  $H_m$ , differs from the Hamiltonian in vacuum  $H_0$ ,  $H_m = H_0 + H_{int}$ , where  $H_{int}$  describes the interaction of neutrinos with the particles of matter. When, for instance,  $\nu_e$  and  $\nu_\mu$  propagate in matter, they can scatter (due to  $H_{int}$ ) on the electrons ( $e^-$ ), protons ( $p$ ) and neutrons ( $n$ ) present in matter. The incoherent elastic and the quasi-elastic scattering, in which the states of the initial particles change in the process (destroying the coherence between the neutrino states), are not of interest - they have a negligible effect on the solar

neutrino propagation in the Sun and on the solar, atmospheric and reactor neutrino propagation in the Earth [144]: even in the center of the Sun, where the matter density is relatively high ( $\sim 150$  g/cm<sup>3</sup>), a  $\nu_e$  with energy of 1 MeV has a mean free path with respect to the indicated scattering processes  $\sim 10^{10}$  km. We recall that the solar radius is much smaller:  $R_\odot = 6.96 \times 10^5$  km. The oscillating  $\nu_e$  and  $\nu_\mu$  can scatter also elastically in the forward direction on the  $e^-$ ,  $p$  and  $n$ , with the momenta and the spin states of the particles remaining unchanged. In such a process the coherence of the neutrino states is preserved.

The  $\nu_e$  and  $\nu_\mu$  coherent elastic scattering on the particles of matter generates nontrivial indices of refraction of the  $\nu_e$  and  $\nu_\mu$  in matter [26]:  $\kappa(\nu_e) \neq 1$ ,  $\kappa(\nu_\mu) \neq 1$ . Most importantly, we have  $\kappa(\nu_e) \neq \kappa(\nu_\mu)$ . The difference  $\kappa(\nu_e) - \kappa(\nu_\mu)$  is determined essentially by the difference of the real parts of the forward  $\nu_e - e^-$  and  $\nu_\mu - e^-$  elastic scattering amplitudes [26]  $\text{Re}[F_{\nu_e-e}(0)] - \text{Re}[F_{\nu_\mu-e}(0)]$ : due to the flavour symmetry of the neutrino - quark (neutrino - nucleon) neutral current interaction, the forward  $\nu_e - p, n$  and  $\nu_\mu - p, n$  elastic scattering amplitudes are equal and therefore do not contribute to the difference of interest [145]. The imaginary parts of the forward scattering amplitudes (responsible, in particular, for decoherence effects) are proportional to the corresponding total scattering cross-sections and in the case of interest are negligible in comparison with the real parts. The real parts of the amplitudes  $F_{\nu_e-e}(0)$  and  $F_{\nu_\mu-e}(0)$  can be calculated in the Standard Model. To leading order in the Fermi constant  $G_F$ , only the term in  $F_{\nu_e-e}(0)$  due to the diagram with exchange of a virtual  $W^\pm$ -boson contributes to  $F_{\nu_e-e}(0) - F_{\nu_\mu-e}(0)$ . One finds the following result for  $\kappa(\nu_e) - \kappa(\nu_\mu)$  in the rest frame of the scatters [26,146,147]:

$$\begin{aligned} \kappa(\nu_e) - \kappa(\nu_\mu) &= \frac{2\pi}{p^2} \left( \text{Re}[F_{\nu_e-e}(0)] - \text{Re}[F_{\nu_\mu-e}(0)] \right) \\ &= -\frac{1}{p} \sqrt{2} G_F N_e, \end{aligned} \quad (14.53)$$

where  $N_e$  is the electron number density in matter. Given  $\kappa(\nu_e) - \kappa(\nu_\mu)$ , the system of evolution equations describing the  $\nu_e \leftrightarrow \nu_\mu$  oscillations in matter reads [26]:

$$i \frac{d}{dt} \begin{pmatrix} A_e(t, t_0) \\ A_\mu(t, t_0) \end{pmatrix} = \begin{pmatrix} -\epsilon(t) & \epsilon' \\ \epsilon' & \epsilon(t) \end{pmatrix} \begin{pmatrix} A_e(t, t_0) \\ A_\mu(t, t_0) \end{pmatrix} \quad (14.54)$$

where  $A_e(t, t_0)$  ( $A_\mu(t, t_0)$ ) is the amplitude of the probability to find  $\nu_e$  ( $\nu_\mu$ ) at time  $t$  of the evolution of the system if at time  $t_0 \leq t$  the neutrino  $\nu_e$  or  $\nu_\mu$  has been produced and

$$\epsilon(t) = \frac{1}{2} \left[ -\frac{\Delta m^2}{2E} \cos 2\theta - \sqrt{2} G_F N_e(t) \right], \quad \epsilon' = \frac{\Delta m^2}{4E} \sin 2\theta. \quad (14.55)$$

The term  $\sqrt{2} G_F N_e(t)$  in  $\epsilon(t)$  accounts for the effects of matter on neutrino oscillations. The system of evolution equations describing the oscillations of antineutrinos  $\bar{\nu}_e \leftrightarrow \bar{\nu}_\mu$  in matter has exactly the same form except for the matter term in  $\epsilon(t)$  which changes sign. The effect of matter in neutrino oscillations is usually called the Mikheyev, Smirnov, Wolfenstein (or MSW) effect.

Consider first the case of  $\nu_e \leftrightarrow \nu_\mu$  oscillations in matter with constant density:  $N_e(t) = N_e = \text{const}$ . Due to the interaction term  $H_{int}$  in  $H_m$ , the eigenstates of the Hamiltonian of the neutrino system in vacuum,  $|\nu_{1,2}\rangle$  are not eigenstates of  $H_m$ . For the eigenstates  $|\nu_{1,2}^m\rangle$  of  $H_m$ , which diagonalize the evolution matrix in the r.h.s. of the system Eq. (14.54) we have:

$$|\nu_e\rangle = |\nu_1^m\rangle \cos \theta_m + |\nu_2^m\rangle \sin \theta_m, \quad |\nu_\mu\rangle = -|\nu_1^m\rangle \sin \theta_m + |\nu_2^m\rangle \cos \theta_m. \quad (14.56)$$

Here  $\theta_m$  is the neutrino mixing angle in matter [26],

$$\sin 2\theta_m = \frac{\tan 2\theta}{\sqrt{(1 - \frac{N_e}{N_e^{\text{res}}})^2 + \tan^2 2\theta}}, \quad \cos 2\theta_m = \frac{1 - N_e/N_e^{\text{res}}}{\sqrt{(1 - \frac{N_e}{N_e^{\text{res}}})^2 + \tan^2 2\theta}}, \quad (14.57)$$

where the quantity

$$N_e^{res} = \frac{\Delta m^2 \cos 2\theta}{2E\sqrt{2}G_F} \cong 6.56 \times 10^6 \frac{\Delta m^2 [\text{eV}^2]}{E [\text{MeV}]} \cos 2\theta \text{ cm}^{-3} N_A, \quad (14.58)$$

is called (for  $\Delta m^2 \cos 2\theta > 0$ ) “resonance density” [27,146],  $N_A$  being Avogadro’s number. The “adiabatic” states  $|\nu_{1,2}^m\rangle$  have energies  $E_{1,2}^m$  whose difference is given by

$$E_2^m - E_1^m = \frac{\Delta m^2}{2E} \left( \left(1 - \frac{N_e}{N_e^{res}}\right)^2 \cos^2 2\theta + \sin^2 2\theta \right)^{\frac{1}{2}} \equiv \frac{\Delta M^2}{2E}. \quad (14.59)$$

The probability of  $\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}$  transition in matter with  $N_e = \text{const.}$  has the form [26,146]

$$P_m^{2\nu}(\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}) = |A_{e(\mu)}(t)|^2 = \frac{1}{2} \sin^2 2\theta_m [1 - \cos 2\pi \frac{L}{L_m}]$$

$$L_m = 2\pi / (E_2^m - E_1^m), \quad (14.60)$$

where  $L_m$  is the oscillation length in matter. As Eq. (14.57) indicates, the dependence of  $\sin^2 2\theta_m$  on  $N_e$  has a resonance character [27]. Indeed, if  $\Delta m^2 \cos 2\theta > 0$ , for any  $\sin^2 2\theta \neq 0$  there exists a value of  $N_e$  given by  $N_e^{res}$ , such that when  $N_e = N_e^{res}$  we have  $\sin^2 2\theta_m = 1$  independently of the value of  $\sin^2 2\theta < 1$ . This implies that the presence of matter can lead to a strong enhancement of the oscillation probability  $P_m^{2\nu}(\nu_{\mu(e)} \rightarrow \nu_{e(\mu)})$  even when the  $\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}$  oscillations in vacuum are suppressed due to a small value of  $\sin^2 2\theta$ . For obvious reasons

$$N_e = N_e^{res} \equiv \frac{\Delta m^2 \cos 2\theta}{2E\sqrt{2}G_F}, \quad (14.61)$$

is called the “resonance condition” [27,146], while the energy at which Eq. (14.61) holds for given  $N_e$  and  $\Delta m^2 \cos 2\theta$ , is referred to as the “resonance energy”,  $E^{res}$ . The oscillation length at resonance is given by [27]  $L_m^{res} = L^v / \sin 2\theta$ , while the width in  $N_e$  of the resonance at half height reads  $\Delta N_e^{res} = 2N_e^{res} \tan 2\theta$ . Thus, if the mixing angle in vacuum is small, the resonance is narrow,  $\Delta N_e^{res} \ll N_e^{res}$ , and  $L_m^{res} \gg L^v$ . The energy difference  $E_2^m - E_1^m$  has a minimum at the resonance:  $(E_2^m - E_1^m)^{res} = \min (E_2^m - E_1^m) = (\Delta m^2 / (2E)) \sin 2\theta$ .

It is instructive to consider two limiting cases. If  $N_e \ll N_e^{res}$ , we have from Eq. (14.57) and Eq. (14.59),  $\theta_m \cong \theta$ ,  $L_m \cong L^v$  and neutrinos oscillate practically as in vacuum. In the limit  $N_e \gg N_e^{res}$ ,  $N_e^{res} \tan^2 2\theta$ , one finds  $\theta_m \cong \pi/2$  ( $\cos 2\theta_m \cong -1$ ) and the presence of matter suppresses the  $\nu_\mu \leftrightarrow \nu_e$  oscillations. In this case  $|\nu_e\rangle \cong |\nu_2^m\rangle$ ,  $|\nu_\mu\rangle = -|\nu_1^m\rangle$ , i.e.,  $\nu_e$  practically coincides with the heavier matter-eigenstate, while  $\nu_\mu$  coincides with the lighter one.

Since the neutral current weak interaction of neutrinos in the Standard Model is flavour symmetric, the formulae and results we have obtained are valid for the case of  $\nu_e - \nu_\tau$  mixing and  $\nu_e \leftrightarrow \nu_\tau$  oscillations in matter as well. The case of  $\nu_\mu - \nu_\tau$  mixing, however, is different: to a relatively good precision we have [148]  $\kappa(\nu_\mu) \cong \kappa(\nu_\tau)$  and the  $\nu_\mu \leftrightarrow \nu_\tau$  oscillations in the matter of the Earth and the Sun proceed practically as in vacuum [149].

The analogs of Eq. (14.57) to Eq. (14.60) for oscillations of antineutrinos,  $\bar{\nu}_\mu \leftrightarrow \bar{\nu}_e$ , in matter can formally be obtained by replacing  $N_e$  with  $(-N_e)$  in the indicated equations. It should be clear that depending on the sign of  $\Delta m^2 \cos 2\theta$ , the presence of matter can lead to resonance enhancement either of the  $\nu_\mu \leftrightarrow \nu_e$  or of the  $\bar{\nu}_\mu \leftrightarrow \bar{\nu}_e$  oscillations, but not of both types of oscillations [146]. For  $\Delta m^2 \cos 2\theta < 0$ , for instance, the matter can only suppress the  $\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}$  oscillations, while it can enhance the  $\bar{\nu}_{\mu(e)} \rightarrow \bar{\nu}_{e(\mu)}$  transitions. The dependence of the effects of matter in  $\nu_\mu \rightarrow \nu_e$  and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations on  $\text{sgn}(\Delta m^2 \cos 2\theta)$  is at basis of the plans to determine the sign of  $\Delta m_{31(32)}^2$ , and thus the type of spectrum neutrino masses obey - with normal or inverted ordering - in long baseline neutrino oscillation experiments (NO $\nu$ A, DUNE) and in atmospheric neutrino experiments with large volume detectors (PINGU, ORCA, INO, Hyper-Kamiokande, DUNE).

The discussed disparity between the behavior of neutrinos and that of antineutrinos is a consequence of the fact that the matter in the

Sun or in the Earth we are interested in is not charge-symmetric (it contains  $e^-$ ,  $p$  and  $n$ , but does not contain their antiparticles) and therefore the oscillations in matter are neither CP- nor CPT-invariant [65]. Thus, even in the case of 2-neutrino mixing and oscillations we have, e.g.,  $P_m^{2\nu}(\nu_\mu \rightarrow \nu_e) \neq P_m^{2\nu}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$  and  $P_m^{2\nu}(\nu_e \rightarrow \nu_{\mu(\tau)}) \neq P_m^{2\nu}(\bar{\nu}_e \rightarrow \bar{\nu}_{\mu(\tau)})$ .

The  $\nu_\mu \leftrightarrow \nu_e$  ( $\bar{\nu}_\mu \leftrightarrow \bar{\nu}_e$ ) and  $\nu_e \leftrightarrow \nu_{\mu(\tau)}$  ( $\bar{\nu}_e \leftrightarrow \bar{\nu}_{\mu(\tau)}$ ) oscillations in matter will be invariant with respect to the operation of time reversal if the  $N_e$  distribution along the neutrino path is symmetric with respect to this operation [63,150]. The latter condition is fulfilled (to a good approximation) for the  $N_e$  distribution along a path of a neutrino crossing the Earth [151].

#### 14.8.1. Effects of Earth matter on oscillations of neutrinos. Analytic expressions for oscillation probabilities :

The formalism we have developed can be applied, e.g., to the study of matter effects in the  $\nu_e \leftrightarrow \nu_{\mu(\tau)}$  ( $\nu_{\mu(\tau)} \leftrightarrow \nu_e$ ) and  $\bar{\nu}_e \leftrightarrow \bar{\nu}_{\mu(\tau)}$  ( $\bar{\nu}_{\mu(\tau)} \leftrightarrow \bar{\nu}_e$ ) oscillations of neutrinos which traverse the Earth [152]. Indeed, the Earth density distribution in the existing Earth models [151] is assumed to be spherically symmetric and there are two major density structures - the core and the mantle, and a certain number of substructures (shells or layers). The Earth radius is  $R_\oplus = 6371$  km; the Earth core has a radius of  $R_c = 3486$  km, so the Earth mantle depth is 2885 km. For a spherically symmetric Earth density distribution, the neutrino trajectory in the Earth is specified by the value of the nadir angle  $\theta_n$  of the trajectory. For  $\theta_n \leq 33.17^\circ$ , or path lengths  $L \geq 10660$  km, neutrinos cross the Earth core. The path length for neutrinos which cross only the Earth mantle is given by  $L = 2R_\oplus \cos \theta_n$ . If neutrinos cross the Earth core, the lengths of the paths in the mantle,  $2L^{\text{man}}$ , and in the core,  $L^{\text{core}}$ , are determined by:  $L^{\text{man}} = R_\oplus \cos \theta_n - (R_c^2 - R_\oplus^2 \sin^2 \theta_n)^{\frac{1}{2}}$ ,  $L^{\text{core}} = 2(R_c^2 - R_\oplus^2 \sin^2 \theta_n)^{\frac{1}{2}}$ . The mean electron number densities in the mantle and in the core according to the PREM model read [151]:  $\bar{N}_e^{\text{man}} \cong 2.2 \text{ cm}^{-3} N_A$ ,  $\bar{N}_e^c \cong 5.4 \text{ cm}^{-3} N_A$ . Thus, we have  $\bar{N}_e^c \cong 2.5 \bar{N}_e^{\text{man}}$ . The change of  $N_e$  from the mantle to the core can well be approximated by a step function [151]. The electron number density  $N_e$  changes relatively little around the indicated mean values along the trajectories of neutrinos which cross a substantial part of the Earth mantle, or the mantle and the core, and the two-layer constant density approximation,  $N_e^{\text{man}} = \text{const.} = \bar{N}_e^{\text{man}}$ ,  $N_e^c = \text{const.} = \bar{N}_e^c$ ,  $\bar{N}_e^{\text{man}}$  and  $\bar{N}_e^c$  being the mean densities along the given neutrino path in the Earth, was shown to be sufficiently accurate in what concerns the calculation of neutrino oscillation probabilities [63,154,155] (and references quoted in [154,155]) in a large number of specific cases. This is related to the fact that the relatively small changes of density along the path of the neutrinos in the mantle (or in the core) take place over path lengths which are typically considerably smaller than the corresponding oscillation length in matter.

In the case of 3-neutrino mixing and for neutrino energies of  $E \gtrsim 2$  GeV, the effects due to  $\Delta m_{21}^2$  ( $|\Delta m_{21}^2| \ll |\Delta m_{31(23)}^2|$ , see Eq. (14.43)) in the neutrino oscillation probabilities are sub-dominant and to leading order can be neglected: the corresponding resonance density  $|N_e^{res}| \lesssim 0.25 \text{ cm}^{-3} N_A \ll \bar{N}_e^{\text{man,c}}$  and the Earth matter strongly suppresses the oscillations due to  $\Delta m_{21}^2$ . For oscillations in vacuum this approximation is valid in the case of NO (IO) neutrino mass spectrum (see Section 2) as long as the leading order contribution due to  $\Delta m_{31(23)}^2$  in the relevant probabilities is bigger than approximately  $10^{-3}$ . In this case the 3-neutrino  $\nu_e \rightarrow \nu_{\mu(\tau)}$  ( $\bar{\nu}_e \rightarrow \bar{\nu}_{\mu(\tau)}$ ) and  $\nu_{\mu(\tau)} \rightarrow \nu_e$  ( $\bar{\nu}_{\mu(\tau)} \rightarrow \bar{\nu}_e$ ) transition probabilities for neutrinos traversing the Earth, reduce effectively to a 2-neutrino transition probability (see, e.g., Refs. [155–157]), with  $\Delta m_{31(23)}^2$  and  $\theta_{13}$  playing the role of the relevant 2-neutrino vacuum oscillation parameters. We note that in the approximation of negligible  $\Delta m_{21}^2$  we have  $\Delta m_{31}^2 = \Delta m_{32}^2$ . Therefore in what follows in this part of the article we will use, whenever relevant, only  $\Delta m_{31}^2$  in the analytic expressions.

As we have discussed in Sections 14.2 and will be discussed in greater detail in Section 14.12, the value of  $\sin^2 2\theta_{13}$  has been determined with a rather high precision in the Daya Bay [38] and RENO [39] experiments. The best fit values found in the two

experiments read, respectively,  $\sin^2 2\theta_{13} = 0.084$  [38] and  $0.087$  [39]. The 3-neutrino oscillation probabilities of the atmospheric and accelerator  $\nu_{e,\mu}$  having energy  $E \gtrsim 2$  GeV and crossing the Earth along a trajectory characterized by a nadir angle  $\theta_n$ , for instance, have the following form in the approximation of negligible  $\Delta m_{21}^2$ :

$$P_m^{3\nu}(\nu_e \rightarrow \nu_e) \cong 1 - P_m^{2\nu}, \quad (14.62)$$

$$P_m^{3\nu}(\nu_e \rightarrow \nu_\mu) \cong P_m^{3\nu}(\nu_\mu \rightarrow \nu_e) \cong s_{23}^2 P_m^{2\nu},$$

$$P_m^{3\nu}(\nu_e \rightarrow \nu_\tau) \cong c_{23}^2 P_m^{2\nu}, \quad (14.63)$$

$$P_m^{3\nu}(\nu_\mu \rightarrow \nu_\mu) \cong 1 - s_{23}^4 P_m^{2\nu} - 2c_{23}^2 s_{23}^2 \left[ 1 - \text{Re} \left( e^{-i\kappa} A_m^{2\nu}(\nu' \rightarrow \nu') \right) \right], \quad (14.64)$$

$$P_m^{3\nu}(\nu_\mu \rightarrow \nu_\tau) = 1 - P_m^{3\nu}(\nu_\mu \rightarrow \nu_\mu) - P_m^{3\nu}(\nu_\mu \rightarrow \nu_e). \quad (14.65)$$

Here  $P_m^{2\nu} \equiv P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}; E, \theta_n)$  is the probability of the 2-neutrino  $\nu_e \rightarrow \nu'$  ( $s_{23}\nu_\mu + c_{23}\nu_\tau$ ) oscillations in the Earth, and  $\kappa$  and  $A_m^{2\nu}(\nu' \rightarrow \nu') \equiv A_m^{2\nu}$  are known phase and 2-neutrino transition probability amplitude (see, e.g., Refs. [155,156]). We note that Eq. (14.62) to Eq. (14.64) are based only on the assumptions that  $|N_{e21}^{res}|$  is much smaller than the densities in the Earth mantle and core and that  $|\Delta m_{21}^2| \ll |\Delta m_{31(23)}^2|$ , and does not rely on the constant density approximation. Similar results are valid for the corresponding antineutrino oscillation probabilities: one has just to replace  $P_m^{2\nu}$ ,  $\kappa$  and  $A_m^{2\nu}$  in the expressions given above with the corresponding quantities for antineutrinos (the latter are obtained from those for neutrinos by changing the sign in front of  $N_e$ ). Obviously, we have:  $P(\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}), P(\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}) \leq \sin^2 2\theta_{23}$ , and  $P(\nu_e \rightarrow \nu_\tau), P(\bar{\nu}_e \rightarrow \bar{\nu}_\tau) \leq \cos^2 2\theta_{23}$ . The one  $\Delta m^2$  dominance approximation and correspondingly Eq. (14.62) to Eq. (14.65) were used by the Super-Kamiokande Collaboration in their 2006 neutrino oscillation analysis of the multi-GeV atmospheric neutrino data [158].

In the case of neutrinos crossing only the Earth mantle and in the constant density approximation,  $P_m^{2\nu}$  is given by the r.h.s. of Eq. (14.60) with  $\theta$ ,  $\Delta m^2$  and  $N_e$  replaced respectively by  $\theta_{13}$ ,  $\Delta m_{31}^2$  and  $\bar{N}_e^{man}$  (corresponding to the given  $\theta_n$ ) in the relevant expressions Eq. (14.57), Eq. (14.58) and Eq. (14.59) for  $\sin 2\theta_m$ ,  $N_e^{res}$  and  $(E_2^m - E_1^m)$ , while for  $\kappa$  and  $A_m^{2\nu}$  we have (see, e.g., Ref. 155):

$$\kappa \cong \frac{1}{2} \left\{ \frac{\Delta m_{31}^2}{2E} L + \sqrt{2} G_F \bar{N}_e^{man} L - \frac{\Delta m_{31}^2 L}{2E} \right\},$$

$$A_m^{2\nu} = 1 + (e^{-i \frac{\Delta m_{31}^2 L}{2E}} - 1) \cos^2 \theta_{13}^m, \quad (14.66)$$

where  $\Delta m_{31}^2$  and  $\theta_{13}^m$  can be obtained from Eq. (14.59) and Eq. (14.57) by setting  $\theta = \theta_{13}$ ,  $\Delta m^2 = \Delta m_{31}^2 > 0$ ,  $N_e^{res} = N_{e31}^{res} = \Delta m_{31}^2 \cos 2\theta_{13} / (2E\sqrt{2}G_F)$  and  $N_e = \bar{N}_e^{man}(\theta_n)$ . Clearly,  $\theta_{13}^m$  is the mixing angle in the mantle which coincides in vacuum with  $\theta_{13}$ . In the expressions for  $P_m^{2\nu} \equiv P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}; E, \theta_n, \bar{N}_e^{man})$ ,  $\kappa$  and  $A_m^{2\nu}$  in the case of oscillations in the mantle,  $L = 2R_\oplus \cos \theta_n$  is the distance the neutrino travels in the mantle. The corresponding expressions for antineutrino oscillations, as we have noticed earlier, can be obtained from those derived above by making the change  $\bar{N}_e^{man} \rightarrow -\bar{N}_e^{man}$ .

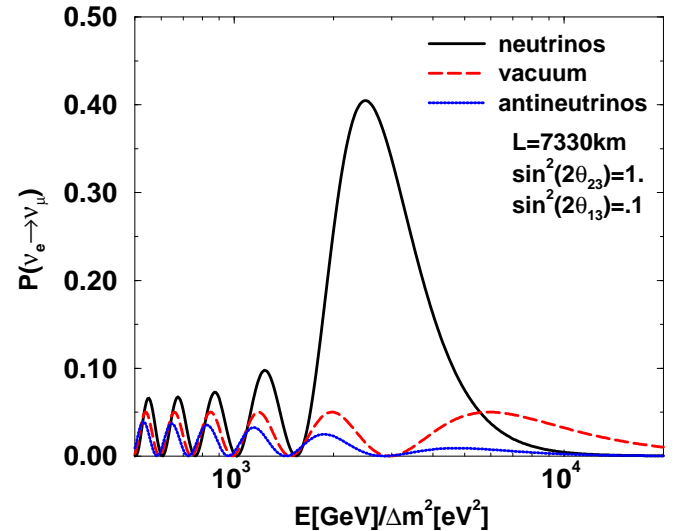
The analytic results for  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}; E, \theta_n, \bar{N}_e^{man})$ ,  $\kappa$  and  $A_m^{2\nu}$ , described above and obtained in the constant mantle density approximation, as we have already remarked, provide a relatively precise description of the  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$ ,  $\nu_e \rightarrow \nu_{\mu(\tau)}$ , etc. oscillation probabilities in the Earth mantle if for each given trajectory of the neutrinos in the mantle, specified by the nadir angle  $\theta_n$ , in the calculations one uses for  $\bar{N}_e^{man}$  the mean value of the electron number density along that specific trajectory:  $\bar{N}_e^{man} = \bar{N}_e^{man}(\theta_n)$ , where  $\bar{N}_e^{man}(\theta_n)$  should be calculated using the density distribution given by the existing Earth models [151].

It follows from Eq. (14.62) and Eq. (14.63) that for  $\Delta m_{31}^2 \cos 2\theta_{13} > 0$ , the oscillation effects of interest, e.g., in the  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  and  $\nu_e \rightarrow \nu_\tau$  transitions will be maximal if  $P_m^{2\nu} \cong 1$ , i.e., if Eq. (14.61) leading to  $\sin^2 2\theta_m \cong 1$  is fulfilled, and ii)  $\cos(\Delta M^2 L / (2E)) \cong -1$ . Given the value of  $\bar{N}_e^{man}$ , the first condition determines the neutrino's energy, while the second determines the path length  $L$ , for which one

can have  $P_m^{2\nu} \cong 1$ . For  $\Delta m_{31}^2 \cong 2.5 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta_{13} \cong 0.090$  and  $\bar{N}_e^{man} \cong 2.2 \text{ N}_A \text{ cm}^{-3}$ , one finds that  $E_{res} \cong 7.1 \text{ GeV}$  and  $L \cong 3522 / \sin 2\theta_{13} \text{ km} \cong 11740 \text{ km}$ . Since for neutrinos crossing only the mantle  $L \lesssim 10660 \text{ km}$ , the second condition can be satisfied only if  $\sin^2 2\theta_{13} \gtrsim 0.11$ , which falls marginally in the  $3\sigma$  range of the experimentally allowed values of  $\sin^2 2\theta_{13}$ . We still get a significant amplification of the probability  $P_m^{2\nu}$ , and therefore of  $P(\nu_{e(\mu)} \rightarrow \nu_{\mu(e)})$  and  $P(\nu_e \rightarrow \nu_\tau)$ , even when  $\cos(\Delta M^2 L / (2E)) = -0.5(-0.2)$ : in this case  $P_m^{2\nu} \cong 0.75$  (0.60). For  $\sin^2 2\theta_{13} \cong 0.090$  we have  $\cos(\Delta M^2 L / (2E)) = -0.5(-0.2)$  if  $L \cong 7826$  (6622) km. Thus, for  $\Delta m_{31}^2 > 0$ , the Earth matter effects can amplify  $P_m^{2\nu}$ , and therefore  $P(\nu_{e(\mu)} \rightarrow \nu_{\mu(e)})$  and  $P(\nu_e \rightarrow \nu_\tau)$ , significantly when the neutrinos cross only the mantle, for  $E \sim 7 \text{ GeV}$  and sufficiently large path lengths  $L$ .

If  $\Delta m_{31}^2 < 0$  the same considerations apply for the corresponding antineutrino oscillation probabilities  $\bar{P}_m^{2\nu} = \bar{P}_m^{2\nu}(\bar{\nu}_e \rightarrow (s_{23}\bar{\nu}_\mu + c_{23}\bar{\nu}_\tau))$  and correspondingly for  $P(\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)})$  and  $P(\bar{\nu}_e \rightarrow \bar{\nu}_\tau)$ . For  $\Delta m_{31}^2 > 0$ , the  $\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}$  and  $\bar{\nu}_e \rightarrow \bar{\nu}_\tau$  oscillations are suppressed by the Earth matter, while if  $\Delta m_{31}^2 < 0$ , the same conclusion holds for the  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  and  $\nu_e \rightarrow \nu_\tau$  oscillations. The dependence on  $\text{sgn}(\Delta m_{31}^2)$  of the effects of Earth matter - enhancement or suppression - on the  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  and  $\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}$  oscillations taking place when the neutrinos traverse the Earth mantle, will be exploited in the current and planned long baseline and atmospheric neutrino oscillation experiments aiming, in particular, to determine the neutrino mass ordering (NO $\nu$ A, DUNE, PINGU, ORCA, INO, Hyper-Kamiokande).

The discussed features of the Earth matter effects in the  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  and  $\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}$  oscillation probabilities for neutrinos with a path length in the Earth mantle of 7330 km and for  $\Delta m_{31}^2 > 0$ ,  $\sin^2 2\theta_{13} = 0.10$  and  $\sin^2 2\theta_{23} = 1$  are illustrated in Fig. 14.7 (taken from Ref. [159]). The amplification of the  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  oscillation probability due to the Earth matter effect in the region of the resonance value of  $E/\Delta m_{31}^2$  and the suppression of the  $\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}$  oscillation probability in the same region are clearly seen in the figure.



**Figure 14.7:** The  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  and  $\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}$  oscillation probabilities given in Eq. (14.63),  $P(\nu_e \rightarrow \nu_\mu) = P(\nu_\mu \rightarrow \nu_e)$  (black solid line) and  $P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu) = P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$  (blue solid line), as functions of  $E/\Delta m^2$  for  $\Delta m^2 \equiv \Delta m_{31}^2 > 0$ ,  $\sin^2 2\theta_{13} = 0.10$  and  $\sin^2 2\theta_{23} = 1$ . The figure is obtained for neutrinos crossing the Earth mantle along a path with length of  $L = 7330 \text{ km}$ . The corresponding vacuum oscillation probability  $P^{vac}(\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}) = P^{vac}(\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)})$  is also shown (red dashed line). For  $\Delta m^2 \equiv \Delta m_{31}^2 < 0$ , the black and blue solid lines will correspond respectively to the probabilities  $P(\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)})$  and  $P(\nu_{e(\mu)} \rightarrow \nu_{\mu(e)})$  (from Ref. 159).

In the case of neutrinos crossing the Earth core, new resonance-like effects become possible in the  $\nu_\mu \rightarrow \nu_e$  and  $\nu_e \rightarrow \nu_{\mu(\tau)}$  (or  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  and  $\bar{\nu}_e \rightarrow \bar{\nu}_{\mu(\tau)}$ ) transitions [154–156,160–162]. For  $\Delta m_{31}^2 > 0$  and certain values of  $\sin^2 \theta_{13} \lesssim 0.05$  we can have [161]  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) \cong 1$ , and correspondingly maximal  $P_m^{3\nu}(\nu_e \rightarrow \nu_\mu) = P_m^{3\nu}(\nu_\mu \rightarrow \nu_e) \cong s_{23}^2$ , only due to the effect of maximal constructive interference between the amplitudes of the  $\nu_e \rightarrow \nu'$  transitions in the Earth mantle and in the Earth core. The effect differs from the MSW one and the enhancement happens in the case of interest at a value of the energy between the MSW resonance energies corresponding to the density in the mantle and that of the core, or at a value of the resonance density  $N_e^{res}$  which lies between the values of  $N_e$  in the mantle and in the core [154]. In Refs. [154,155] the enhancement was called “neutrino oscillation length resonance (NOLR)”, while in Refs. [156,160] the term “parametric resonance” for the same effect was used [163]. The *mantle-core enhancement effect (or NOLR)* is caused by the existence (for a given neutrino trajectory through the Earth core) of points of resonance-like maximal neutrino conversion,  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) = 1$ , in the corresponding space of neutrino oscillation parameters [161]. For  $\Delta m_{31}^2 < 0$  the mantle-core enhancement can take place for the antineutrino transitions,  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  and  $\bar{\nu}_e \rightarrow \bar{\nu}_{\mu(\tau)}$ . For neutrinos crossing the Earth core, analytic expressions for  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13})$  and  $\kappa, A_m^{2\nu}$  were derived in the two-layer constant density approximation for the Earth density distribution in [154] and [155], respectively.

A rather complete set of values of  $\Delta m_{31}^2/E > 0$  and  $\sin^2 2\theta_{13}$  for which  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) = 1$  was found in Ref. 161. In the two-layer constant density approximation, the values of  $\Delta m_{31}^2/E > 0$  and  $\sin^2 2\theta_{13}$  at which  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) = 1$  can be derived as solutions of the following system of equations [161]:

$$\tan \frac{\phi^{man}}{2} = \pm \sqrt{\frac{-\cos 2\theta_{13}^{core}}{\cos(2\theta_{13}^{core} - 4\theta_{13}^{man})}}, \quad (14.67)$$

$$\tan \frac{\phi^{core}}{2} = \pm \sqrt{\frac{\cos 2\theta_{13}^{man}}{-\cos(2\theta_{13}^{core}) \cos(2\theta_{13}^{core} - 4\theta_{13}^{man})}}, \quad (14.68)$$

where the signs in the two equations are correlated,  $\phi^{man} = (E_3^m - E_1^m)_{man} 2L^{man}$ ,  $\phi^{core} = (E_3^m - E_1^m)_{core} 2L^{core}$ , and  $L^{core}$  are the neutrino path lengths in the Earth mantle and the core, and  $\theta_{13}^{man}$  and  $\theta_{13}^{core}$  are the values of the angle  $\theta_{13}$  in the mantle and in the core. The expressions for  $(E_3^m - E_1^m)_{man}$  and  $(E_3^m - E_1^m)_{core}$  can be obtained respectively from Eq. (14.59) and Eq. (14.57) by setting  $\theta = \theta_{13}$ ,  $\Delta m^2 = \Delta m_{31}^2$ ,  $N_e^{res} = N_e^{res} = \Delta m_{31}^2 \cos 2\theta_{13} / (2E\sqrt{2}G_F)$  and  $N_e = \tilde{N}_e^{man}(\theta_n)$  ( $N_e = \tilde{N}_e^{core}(\theta_n)$ ).

The location of the points where  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) = 1$  in the  $\Delta m_{31}^2/E - \sin^2 2\theta_{13}$  plane determines the regions in the plane where  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13})$  is large,  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) \gtrsim 0.5$ . These regions vary slowly with the nadir angle, being remarkably wide in the nadir angle and rather wide in the neutrino energy [161], so that the transitions of interest can produce noticeable effects in the measured observables. For  $\sin^2 \theta_{13} \lesssim 0.05$ , there are two sets of values of  $(\Delta m_{31}^2/E, \sin^2 \theta_{13})$  for which  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) = 1$ , and thus two regions in  $\Delta m_{31}^2/E - \sin^2 2\theta_{13}$  plane where  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) \gtrsim 0.5$ . For  $\Delta m_{31}^2 = 2.5 \times 10^{-3}$  eV<sup>2</sup> and nadir angle, e.g.,  $\theta_n = 0$  (Earth center crossing neutrinos), we have  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) = 1$  at  $(E, \sin^2 2\theta_{13}) = (3.4 \text{ GeV}, 0.034)$  and  $(5.2 \text{ GeV}, 0.15)$ . At the same time for  $E = 3.4 \text{ GeV}$  (5.2 GeV), the probability  $P_m^{2\nu}(\Delta m_{31}^2, \theta_{13}) \gtrsim 0.5$  for the values of  $\sin^2 2\theta_{13}$  from the interval  $0.02 \lesssim \sin^2 2\theta_{13} \lesssim 0.10$  ( $0.04 \lesssim \sin^2 2\theta_{13} \lesssim 0.26$ ). Similar results hold for neutrinos crossing the Earth core along the trajectories with  $\theta_n \neq 0$  (for further details see the last article in Ref. 161; see also the last article in Ref. 162).

The mantle-core enhancement of  $P_m^{2\nu}$  (or  $\bar{P}_m^{2\nu}$ ) is relevant, in particular, for the searches of sub-dominant  $\nu_{e(\mu)} \rightarrow \nu_{\mu(e)}$  (or  $\bar{\nu}_{e(\mu)} \rightarrow \bar{\nu}_{\mu(e)}$ ) oscillations of atmospheric neutrinos having energies  $E \gtrsim 2 \text{ GeV}$  and crossing the Earth core on the way to the detector (see Ref. 154 to Ref. 162 and the references quoted therein).

The effects of Earth matter on the oscillations of atmospheric and accelerator neutrinos have not been observed so far. At present there

are no compelling evidences for oscillations of the atmospheric  $\nu_e$  and/or  $\bar{\nu}_e$ .

In the case of oscillations of atmospheric neutrinos in the Earth one has to take into account also the following considerations. The fluxes of atmospheric  $\nu_{e,\mu}$  of energy  $E$ , which reach the detector after crossing the Earth along a given trajectory specified by the value of  $\theta_n$ ,  $\Phi_{\nu_{e,\mu}}(E, \theta_n)$ , are given by the following expressions in the case of the 3-neutrino oscillations under discussion [155,156]:

$$\Phi_{\nu_e}(E, \theta_n) \cong \Phi_{\nu_e}^0 \left( 1 + [s_{23}^2 r - 1] P_m^{2\nu} \right), \quad (14.69)$$

$$\Phi_{\nu_\mu}(E, \theta_n) \cong \Phi_{\nu_\mu}^0 \left( 1 + s_{23}^4 [(s_{23}^2 r)^{-1} - 1] P_m^{2\nu} - 2c_{23}^2 s_{23}^2 [1 - \text{Re}(e^{-i\kappa} A_m^{2\nu}(\nu_\tau \rightarrow \nu_\tau)) \right], \quad (14.70)$$

where  $\Phi_{\nu_{e(\mu)}}^0 = \Phi_{\nu_{e(\mu)}}^0(E, \theta_n)$  is the  $\nu_{e(\mu)}$  flux in the absence of neutrino oscillations and

$$r \equiv r(E, \theta_n) \equiv \frac{\Phi_{\nu_\mu}^0(E, \theta_n)}{\Phi_{\nu_e}^0(E, \theta_n)}. \quad (14.71)$$

It follows from the global analyses of the neutrino oscillation data that the neutrino mixing parameter  $s_{23}^2$  lies (at  $3\sigma$  CL) in the interval (0.38 - 0.64). For NO (IO) neutrino mass spectrum, the three groups which performed recent global analyses, obtained the following best fit values of  $s_{23}^2$  (see Section 14.2): 0.452 (0.455) [52], 0.452 (0.579) [53] and 0.567 (0.573) [54], while in the latest analysis [60] the authors find 0.437 (0.569). For the predicted ratio  $r(E, \theta_n)$  of the atmospheric  $\nu_\mu$  and  $\nu_e$  fluxes for i) the Earth core crossing and ii) only mantle crossing neutrinos, having trajectories for which  $0.3 \lesssim \cos \theta_n \leq 1.0$ , one has [164]  $r(E, \theta_n) \cong (2.6 \div 4.5)$  for neutrinos giving the main contribution to the multi-GeV samples,  $E \cong (2 \div 10) \text{ GeV}$ . Thus, for  $s_{23}^2 = 0.5$  (0.64) one finds for the multi-GeV neutrinos:  $s_{23}^4 [1 - (s_{23}^2 r(E, \theta_z))^{-1}] \cong 0.06 - 0.14$  (0.16 - 0.27) and  $(s_{23}^2 r(E, \theta_z) - 1) \cong 0.3 - 1.3$  (0.66 - 1.9). Thus, the impact of the possible enhancement of  $P_m^{2\nu} \cong 1$  would be largest for the flux of multi-GeV  $\nu_e$ ,  $\Phi_{\nu_e}(E, \theta_n)$ , traversing the Earth. As the preceding discussion suggests and detailed calculations show (see the first two articles quoted in Ref. 162), the sensitivity of the atmospheric neutrino experiments to the neutrino mass ordering depends strongly on the chosen value of  $\sin^2 \theta_{23}$  from its  $3\sigma$  allowed range: it is maximal (minimal) for the maximal (minimal) allowed value of  $\sin^2 \theta_{23}$ . In the case of the planned Hyper-Kamiokande detector and neutrino mass spectrum with normal ordering, for instance, it is estimated [165] that for  $\sin^2 \theta_{23} = 0.60$  it will take approximately (2-3) years to establish at  $3\sigma$  CL that the spectrum is of the NO type, and approximately 10 years if  $\sin^2 \theta_{23} = 0.40$ .

For water Cerenkov detectors, the charged current (CC)  $\nu_l - N$  interaction cross section for multi-GeV neutrinos is approximately by a factor of 2 bigger than the  $\bar{\nu}_l - N$  CC interaction cross section. Since these detectors do not distinguish between the neutrino and anti-neutrino induced CC events, determining that the neutrino mass spectrum is with inverted ordering would require roughly by a factor of 2 longer period of data acquisition than if the spectrum were with normal ordering.

The effects under discussion are larger, in general, for the multi-GeV neutrinos than for the sub-GeV neutrinos having energies  $E \cong (0.1 - 1.0) \text{ GeV}$ . Indeed, for the sub-GeV  $\nu_e$  flux one finds in the limit of negligible  $\theta_{13}$  [166]:  $\Phi_{\nu_e}(E, \theta_n) \cong \Phi_{\nu_e}^0 (1 + [c_{23}^2 r - 1] \bar{P}_m^{2\nu})$ , where  $\bar{P}_m^{2\nu} \equiv \bar{P}_m^{2\nu}(\Delta m_{21}^2, \theta_{12}; E, \theta_n)$  is the probability of the 2-neutrino oscillations in the Earth due to  $\Delta m_{21}^2$  and 2-neutrino mixing with angle  $\theta_{12}$ . For the neutrinos giving contribution to the sub-GeV samples of Super-Kamiokande events one has [164]  $r(E, \theta_z) \cong 2.0$ . If  $s_{23}^2 = 0.5$  and  $r(E, \theta_z) \cong 2.0$ , we get  $(c_{23}^2 r(E, \theta_z) - 1) \cong 0$ , and the possible effects of the  $\nu_\mu \rightarrow \nu_e$  and  $\nu_e \rightarrow \nu_{\mu(\tau)}$  transitions on the  $\nu_e$  flux, and correspondingly in the sub-GeV  $e$ -like samples of events, would be rather strongly suppressed independently of the values of the corresponding transition probabilities.

The same conclusions are valid for the effects of oscillations on the fluxes of, and event rates due to, atmospheric antineutrinos  $\bar{\nu}_e$  and  $\bar{\nu}_\mu$ .

The formulae for anti-neutrino fluxes and oscillation probabilities are analogous to those for neutrinos (see, *e.g.*, Refs. [155,156,162,166]).

The expression for the probability of the  $\nu_\mu \rightarrow \nu_e$  oscillations taking place in the Earth mantle in the case of 3-neutrino mixing, in which both neutrino mass squared differences  $\Delta m_{21}^2$  and  $\Delta m_{31}^2$  contribute and the CP violation effects due to the Dirac phase in the neutrino mixing matrix are taken into account, has the following form in the constant density approximation and keeping terms up to second order in the two small parameters  $|\alpha| \equiv |\Delta m_{21}^2|/|\Delta m_{31}^2| \ll 1$  and  $\sin^2 \theta_{13} \ll 1$  [167]:

$$P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e) \cong P_0 + P_{\sin \delta} + P_{\cos \delta} + P_3. \quad (14.72)$$

Here

$$P_0 = \sin^2 \theta_{23} \frac{\sin^2 2\theta_{13}}{(A-1)^2} \sin^2[(A-1)\Delta] \\ P_3 = \alpha^2 \cos^2 \theta_{23} \frac{\sin^2 2\theta_{12}}{A^2} \sin^2(A\Delta), \quad (14.73)$$

$$P_{\sin \delta} = -\alpha \frac{8J_{CP}}{A(1-A)} (\sin \Delta) (\sin A\Delta) (\sin[(1-A)\Delta]), \quad (14.74)$$

$$P_{\cos \delta} = \alpha \frac{8J_{CP} \cot \delta}{A(1-A)} (\cos \Delta) (\sin A\Delta) (\sin[(1-A)\Delta]), \quad (14.75)$$

where

$$\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}, \quad \Delta = \frac{\Delta m_{31}^2 L}{4E}, \quad A = \sqrt{2} G_F N_e^{\text{man}} \frac{2E}{\Delta m_{31}^2}, \quad (14.76)$$

$\cot \delta = J_{CP}^{-1} \text{Re}(U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*) \propto \cos \delta$ , and we recall that  $J_{CP} = \text{Im}(U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*)$ . The analytic expression for  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  given above is valid for [167] neutrino path lengths in the mantle ( $L \leq 10660$  km) satisfying  $L \lesssim 10560$  km  $E[\text{GeV}] (7.6 \times 10^{-5} \text{ eV}^2/\Delta m_{21}^2)$ , and energies  $E \gtrsim 0.34$  GeV ( $\Delta m_{21}^2/7.6 \times 10^{-5} \text{ eV}^2$ ) ( $1.4 \text{ cm}^{-3} N_A/N_e^{\text{man}}$ ). The expression for the  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation probability can be obtained formally from that for  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  by making the changes  $A \rightarrow -A$  and  $J_{CP} \rightarrow -J_{CP}$ , with  $J_{CP} \cot \delta \equiv \text{Re}(U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*)$  remaining unchanged. The term  $P_{\sin \delta}$  in  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  would be equal to zero if the Dirac phase in the neutrino mixing matrix  $U$  possesses a CP-conserving value. Even in this case, however, we have  $A_{CP}^{(\mu)} \text{ man} \equiv (P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e) - P_m^{3\nu\text{ man}}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)) \neq 0$  due to the effects of the Earth matter. It will be important to experimentally disentangle the effects of the Earth matter and of  $J_{CP}$  in  $A_{CP}^{(\mu)} \text{ man}$ : this will allow to get information about the Dirac CP violation phase in  $U$ . This can be done, in principle, by studying the energy dependence of  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  and  $P_m^{3\nu\text{ man}}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ . Since the sign of  $\Delta m_{31(32)}^2$  determines for given  $L$  whether the probability  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  or  $P_m^{3\nu\text{ man}}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ , as a function of energy, can be resonantly enhanced or suppressed by the matter effects, the study of the energy dependence of  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  and/or of  $P_m^{3\nu\text{ man}}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$  can provide also information on  $\text{sgn}(\Delta m_{31(32)}^2)$ . In the vacuum limit of  $N_e^{\text{man}} = 0$  ( $A = 0$ ) we have  $A_{CP}^{(\mu)} \text{ man} = A_{CP}^{(\mu)}$  (see Eq. (14.40)) and only the term  $P_{\sin \delta}$  contributes to the asymmetry  $A_{CP}^{(\mu)}$ .

The preceding remarks apply also to the probabilities  $P_m^{3\nu\text{ man}}(\nu_e \rightarrow \nu_\mu)$  and  $P_m^{3\nu\text{ man}}(\bar{\nu}_e \rightarrow \bar{\nu}_\mu)$ . The probability  $P_m^{3\nu\text{ man}}(\nu_e \rightarrow \nu_\mu)$ , for example, can formally be obtained from the expression for the probability  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  by changing the sign of the term  $P_{\sin \delta}$ .

The expression for the probability  $P_m^{3\nu\text{ man}}(\nu_\mu \rightarrow \nu_e)$  given in Eq. (14.72) and the corresponding expression for  $P_m^{3\nu\text{ man}}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$  can be used for the interpretation of the data of the current and the future planned long baseline oscillation experiments T2K, NO $\nu$ A, MINOS+, DUNE [84] and T2HK [85].

## 14.8.2. Oscillations of solar neutrinos :

### 14.8.2.1. Qualitative analysis:

Consider next the oscillations of solar  $\nu_e$  while they propagate from the central part of the Sun, where they are produced, to the surface of the Sun [27,153] (see also Ref. 26 and, *e.g.*, Ref. 168). Details concerning the production, spectrum, magnitude and particularities of the solar neutrino flux were discussed in Section 14.6, while the methods of detection of solar neutrinos, description of solar neutrino experiments and of the data they provided will be discussed in the next section (see also Ref. 169). The electron number density  $N_e$  changes considerably along the neutrino path in the Sun: it decreases monotonically from the value of  $\sim 100 \text{ cm}^{-3}$   $N_A$  in the center of the Sun to 0 at the surface of the Sun. According to the contemporary solar models (see, *e.g.*, Ref. [169,99]),  $N_e$  decreases approximately exponentially in the radial direction towards the surface of the Sun:

$$N_e(t) = N_e(t_0) \exp \left\{ -\frac{t-t_0}{r_0} \right\}, \quad (14.77)$$

where  $(t-t_0) \cong d$  is the distance traveled by the neutrino in the Sun,  $N_e(t_0)$  is the electron number density at the point of  $\nu_e$  production in the Sun,  $r_0$  is the scale-height of the change of  $N_e(t)$  and one has [169,99]  $r_0 \sim 0.1 R_\odot$ .

Consider the case of 2-neutrino mixing, Eq. (14.56). Obviously, if  $N_e$  changes with  $t$  (or equivalently with the distance) along the neutrino trajectory, the matter-eigenstates, their energies, the mixing angle and the oscillation length in matter, become, through their dependence on  $N_e$ , also functions of  $t$ :  $|\nu_{1,2}^m\rangle = |\nu_{1,2}^m(t)\rangle$ ,  $E_{1,2}^m = E_{1,2}^m(t)$ ,  $\theta_m = \theta_m(t)$  and  $L_m = L_m(t)$ . It is not difficult to understand qualitatively the possible behavior of the neutrino system when solar neutrinos propagate from the center to the surface of the Sun if one realizes that one is dealing effectively with a two-level system whose Hamiltonian depends on time and admits “jumps” from one level to the other (see Eq. (14.54)). Consider the case of  $\Delta m^2 \cos 2\theta > 0$ . Let us assume first for simplicity that the electron number density at the point of a solar  $\nu_e$  production in the Sun is much bigger than the resonance density,  $N_e(t_0) \gg N_e^{\text{res}}$ . Actually, this is one of the cases relevant to the solar neutrinos. In this case we have  $\theta_m(t_0) \cong \pi/2$  and the state of the electron neutrino in the initial moment of the evolution of the system practically coincides with the heavier of the two matter-eigenstates:

$$|\nu_e\rangle \cong |\nu_2^m(t_0)\rangle. \quad (14.78)$$

Thus, at  $t_0$  the neutrino system is in a state corresponding to the “level” with energy  $E_2^m(t_0)$ . When neutrinos propagate to the surface of the Sun they cross a layer of matter in which  $N_e = N_e^{\text{res}}$ : in this layer the difference between the energies of the two “levels” ( $E_2^m(t) - E_1^m(t)$ ) has a minimal value on the neutrino trajectory (Eq. (14.59) and Eq. (14.61)). Correspondingly, the evolution of the neutrino system can proceed basically in two ways. First, the system can stay on the “level” with energy  $E_2^m(t)$ , *i.e.*, can continue to be in the state  $|\nu_2^m(t)\rangle$  up to the final moment  $t_s$ , when the neutrino reaches the surface of the Sun. At the surface of the Sun  $N_e(t_s) = 0$  and therefore  $\theta_m(t_s) = \theta$ ,  $|\nu_{1,2}^m(t_s)\rangle \equiv |\nu_{1,2}\rangle$  and  $E_{1,2}^m(t_s) = E_{1,2}$ . Thus, in this case the state describing the neutrino system at  $t_0$  will evolve continuously into the state  $|\nu_2\rangle$  at the surface of the Sun. Using Eq. (14.51) with  $l = e$  and  $x = \mu$ , it is easy to obtain the probabilities to find  $\nu_e$  and  $\nu_\mu$  at the surface of the Sun:

$$P(\nu_e \rightarrow \nu_e; t_s, t_0) \cong |\langle \nu_e | \nu_2 \rangle|^2 = \sin^2 \theta \\ P(\nu_e \rightarrow \nu_\mu; t_s, t_0) \cong |\langle \nu_\mu | \nu_2 \rangle|^2 = \cos^2 \theta. \quad (14.79)$$

It is clear that under the assumption made and if  $\sin^2 \theta \ll 1$ , practically a total  $\nu_e \rightarrow \nu_\mu$  conversion is possible. This type of evolution of the neutrino system and the  $\nu_e \rightarrow \nu_\mu$  transitions taking place during the evolution, are called [27] “adiabatic.” They are characterized by the fact that the probability of the “jump” from the upper “level” (having energy  $E_2^m(t)$ ) to the lower “level” (with energy  $E_1^m(t)$ ),  $P'$ , or equivalently the probability of the  $\nu_2^m(t_0) \rightarrow \nu_1^m(t_s)$



transition,  $P' \equiv P'(\nu_2^m(t_0) \rightarrow \nu_1^m(t_s))$ , on the whole neutrino trajectory is negligible:

$$P' \equiv P'(\nu_2^m(t_0) \rightarrow \nu_1^m(t_s)) \cong 0 : \text{adiabatic transitions.} \quad (14.80)$$

The second possibility is realized if in the resonance region, where the two “levels” approach each other most, the system “jumps” from the upper “level” to the lower “level” and after that continues to be in the state  $|\nu_1^m(t)\rangle$  until the neutrino reaches the surface of the Sun. Evidently, now we have  $P' \equiv P'(\nu_2^m(t_0) \rightarrow \nu_1^m(t_s)) \sim 1$ . In this case the neutrino system ends up in the state  $|\nu_1^m(t_s)\rangle \equiv |\nu_1\rangle$  at the surface of the Sun and

$$\begin{aligned} P(\nu_e \rightarrow \nu_e; t_s, t_0) &\cong |\langle \nu_e | \nu_1 \rangle|^2 = \cos^2 \theta \\ P(\nu_e \rightarrow \nu_\mu; t_s, t_0) &\cong |\langle \nu_\mu | \nu_1 \rangle|^2 = \sin^2 \theta. \end{aligned} \quad (14.81)$$

Obviously, if  $\sin^2 \theta \ll 1$ , practically no transitions of the solar  $\nu_e$  into  $\nu_\mu$  will occur. The considered regime of evolution of the neutrino system and the corresponding  $\nu_e \rightarrow \nu_\mu$  transitions are usually referred to as “extremely nonadiabatic.”

Clearly, the value of the “jump” probability  $P'$  plays a crucial role in the  $\nu_e \rightarrow \nu_\mu$  transitions: it fixes the type of the transition and determines to a large extent the  $\nu_e \rightarrow \nu_\mu$  transition probability [153,170,171]. We have considered above two limiting cases. Obviously, there exists a whole spectrum of possibilities since  $P'$  can have any value from 0 to  $\cos^2 \theta$  [172,173]. In general, the transitions are called “nonadiabatic” if  $P'$  is non-negligible.

Numerical studies have shown [27] that solar neutrinos can undergo both adiabatic and nonadiabatic  $\nu_e \rightarrow \nu_\mu$  transitions in the Sun and the matter effects can be substantial in the solar neutrino oscillations for  $10^{-8} \text{ eV}^2 \lesssim \Delta m^2 \lesssim 10^{-4} \text{ eV}^2$ ,  $10^{-4} \lesssim \sin^2 2\theta < 1.0$ .

The condition of adiabaticity of the solar  $\nu_e$  transitions in Sun can be written as [153,170]

$$\begin{aligned} \gamma(t) &\equiv \sqrt{2} G_F \frac{(N_e^{\text{res}})^2}{|\dot{N}_e(t)|} \tan^2 2\theta \left(1 + \tan^{-2} 2\theta_m(t)\right)^{\frac{3}{2}} \gg 1 \\ &\text{adiabatic transitions,} \end{aligned} \quad (14.82)$$

while if  $\gamma(t) \lesssim 1$  the transitions are nonadiabatic (see also Ref. 173), where  $\dot{N}_e(t) \equiv \frac{d}{dt} N_e(t)$ . Condition in Eq. (14.82) implies that the  $\nu_e \rightarrow \nu_{\mu(\tau)}$  transitions in the Sun will be adiabatic if  $N_e(t)$  changes sufficiently slowly along the neutrino path. In order for the transitions to be adiabatic, condition in Eq. (14.82) has to be fulfilled at any point of the neutrino’s path in the Sun.

#### 14.8.2.2. The solar $\nu_e$ survival probability:

The system of evolution equations Eq. (14.54) can be solved exactly for  $N_e$  changing exponentially, Eq. (14.77), along the neutrino path in the Sun [172,174]. More specifically, the system in Eq. (14.54) is equivalent to one second order differential equation (with appropriate initial conditions). The latter can be shown [175] to coincide in form, in the case of  $N_e$  given by Eq. (14.77), with the Schroedinger equation for the radial part of the nonrelativistic wave function of the Hydrogen atom [176]. On the basis of the exact solution, which is expressed in terms of confluent hypergeometric functions, it was possible to derive a complete, simple and very accurate analytic description of the matter-enhanced transitions of solar neutrinos in the Sun for any values of  $\Delta m^2$  and  $\theta$  [26,172,173,177,178] (see also Refs. [27,153,171,179,180]).

The probability that a  $\nu_e$ , produced at time  $t_0$  in the central part of the Sun, will not transform into  $\nu_{\mu(\tau)}$  on its way to the surface of the Sun (reached at time  $t_s$ ) is given by

$$P_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0) = \bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0) + \text{Oscillating terms.} \quad (14.83)$$

Here

$$\bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0) \equiv \bar{P}_{\odot} = \frac{1}{2} + \left(\frac{1}{2} - P'\right) \cos 2\theta_m(t_0) \cos 2\theta, \quad (14.84)$$

is the average survival probability for  $\nu_e$  having energy  $E \cong p$  [171], where

$$P' = \frac{\exp\left[-2\pi r_0 \frac{\Delta m^2}{2E} \sin^2 \theta\right] - \exp\left[-2\pi r_0 \frac{\Delta m^2}{2E}\right]}{1 - \exp\left[-2\pi r_0 \frac{\Delta m^2}{2E}\right]}, \quad (14.85)$$

is [172] the “jump” probability for exponentially varying  $N_e$ , and  $\theta_m(t_0)$  is the mixing angle in matter at the point of  $\nu_e$  production [179]. The expression for  $\bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0)$  with  $P'$  given by Eq. (14.85) is valid for  $\Delta m^2 > 0$ , but for both signs of  $\cos 2\theta \neq 0$  [172,180]; it is valid for any given value of the distance along the neutrino trajectory and does not take into account the finite dimensions of the region of  $\nu_e$  production in the Sun. This can be done by integrating over the different neutrino paths, *i.e.*, over the region of  $\nu_e$  production.

The oscillating terms in the probability  $P_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0)$  [177,175] were shown [178] to be strongly suppressed for  $\Delta m^2 \gtrsim 10^{-7} \text{ eV}^2$  by the various averagings one has to perform when analyzing the solar neutrino data. The current solar neutrino and KamLAND data suggest that  $\Delta m^2 \cong 7.4 \times 10^{-5} \text{ eV}^2$ . For  $\Delta m^2 \gtrsim 10^{-7} \text{ eV}^2$ , the averaging over the region of neutrino production in the Sun *etc.* renders negligible all interference terms which appear in the probability of  $\nu_e$  survival due to the  $\nu_e \leftrightarrow \nu_{\mu(\tau)}$  oscillations in vacuum taking place on the way of the neutrinos from the surface of the Sun to the surface of the Earth. Thus, the probability that  $\nu_e$  will remain  $\nu_e$  while it travels from the central part of the Sun to the surface of the Earth is effectively equal to the probability of survival of the  $\nu_e$  while it propagates from the central part to the surface of the Sun and is given by the average probability  $\bar{P}_{\odot}(\nu_e \rightarrow \nu_e; t_s, t_0)$  (determined by Eq. (14.84) and Eq. (14.85)).

If the solar  $\nu_e$  transitions are adiabatic ( $P' \cong 0$ ) and  $\cos 2\theta_m(t_0) \cong -1$  (*i.e.*,  $N_e(t_0)/|N_e^{\text{res}}| \gg 1$ ,  $|\tan 2\theta|$ , the  $\nu_e$  are born “above” (in  $N_e$ ) the resonance region), one has [27]

$$\bar{P}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0) \cong \frac{1}{2} - \frac{1}{2} \cos 2\theta. \quad (14.86)$$

The regime under discussion is realized for  $\sin^2 2\theta \cong 0.8$  (suggested by the data, Sections 14.2 and 14.9.1), if  $E/\Delta m^2$  lies approximately in the range  $(2 \times 10^4 - 3 \times 10^7) \text{ MeV/eV}^2$  (see Ref. 173). This result is relevant for the interpretation of the Super-Kamiokande and SNO solar neutrino data. We see that depending on the sign of  $\cos 2\theta \neq 0$ ,  $\bar{P}^{2\nu}(\nu_e \rightarrow \nu_e)$  is either bigger or smaller than 1/2. It follows from the solar neutrino data that in the range of validity (in  $E/\Delta m^2$ ) of Eq. (14.86) we have  $\bar{P}^{2\nu}(\nu_e \rightarrow \nu_e) \cong 0.3$ . Thus, the possibility of  $\cos 2\theta \leq 0$  is ruled out by the data. Given the choice  $\Delta m^2 > 0$  we made, the data imply that  $\Delta m^2 \cos 2\theta > 0$ .

If  $E/\Delta m^2$  is sufficiently small so that  $N_e(t_0)/|N_e^{\text{res}}| \ll 1$ , we have  $P' \cong 0$ ,  $\theta_m(t_0) \cong \theta$  and the oscillations take place in the Sun as in vacuum [27]:

$$\bar{P}^{2\nu}(\nu_e \rightarrow \nu_e; t_s, t_0) \cong 1 - \frac{1}{2} \sin^2 2\theta, \quad (14.87)$$

which is the average two-neutrino vacuum oscillation probability. This expression describes with good precision the transitions of the solar  $pp$  neutrinos (Section 14.9.1). The extremely nonadiabatic  $\nu_e$  transitions in the Sun, characterized by  $\gamma(t) \ll 1$ , are also described by the average vacuum oscillation probability (Eq. (14.87)) (for  $\Delta m^2 \cos 2\theta > 0$  in this case we have (see *e.g.*, Refs. [172,173])  $\cos 2\theta_m(t_0) \cong -1$  and  $P' \cong \cos^2 \theta$ ).

The probability of  $\nu_e$  survival in the case 3-neutrino mixing takes a simple form for  $|\Delta m_{31}^2| \cong 2.5 \times 10^{-3} \text{ eV}^2 \gg |\Delta m_{21}^2|$ . Indeed, for the energies of solar neutrinos  $E \lesssim 10 \text{ MeV}$ ,  $N_e^{\text{res}}$  corresponding to  $|\Delta m_{31}^2|$  satisfies  $N_e^{\text{res}} \gtrsim 10^3 \text{ cm}^{-3} N_A$  and is by a factor of 10 bigger than  $N_e$  in the center of the Sun. As a consequence, the oscillations due to  $\Delta m_{31}^2$  proceed as in vacuum. The oscillation length associated with  $|\Delta m_{31}^2|$  satisfies  $L_{31}^{\nu} \lesssim 10 \text{ km} \ll \Delta R$ ,  $\Delta R$  being the dimension of the region of  $\nu_e$  production in the Sun. We have for the different components of the solar  $\nu_e$  flux [169]  $\Delta R \cong (0.04 - 0.20) R_{\odot}$ . Therefore the averaging over  $\Delta R$  strongly suppresses the oscillations due to  $\Delta m_{31}^2$  and we get [157,181]:

$$P_{\odot}^{3\nu} \cong \sin^4 \theta_{13} + \cos^4 \theta_{13} P_{\odot}^{2\nu}(\Delta m_{21}^2, \theta_{12}; N_e \cos^2 \theta_{13}), \quad (14.88)$$

where  $P_{\odot}^{2\nu}(\Delta m_{21}^2, \theta_{12}; N_e \cos^2 \theta_{13})$  is given by Eq. (14.83) to Eq. (14.85) in which  $\Delta m^2 = \Delta m_{21}^2$ ,  $\theta = \theta_{12}$  and the solar  $e^-$  number density  $N_e$  is replaced by  $N_e \cos^2 \theta_{13}$ . Thus, the solar  $\nu_e$  transitions observed by the Super-Kamiokande and SNO experiments are described approximately by:

$$P_{\odot}^{3\nu} \cong \sin^4 \theta_{13} + \cos^4 \theta_{13} \sin^2 \theta_{12}. \quad (14.89)$$

The data show that  $P_{\odot}^{3\nu} \cong 0.3$ , which is a strong evidence for matter effects in the solar  $\nu_e$  transitions [182] since in the case of oscillations in vacuum  $P_{\odot}^{3\nu} \cong \sin^4 \theta_{13} + (1 - 0.5 \sin^2 2\theta_{12}) \cos^4 \theta_{13} \gtrsim 0.53$ , where we have used  $\sin^2 \theta_{13} \lesssim 0.0246$  and  $\sin^2 2\theta_{12} \lesssim 0.915$  (see Section 14.2).

The analytic expression for the solar  $\nu_e$  survival probability, Eq. (14.88), with  $P_{\odot}^{2\nu}(\Delta m_{21}^2, \theta_{12}; N_e \cos^2 \theta_{13})$  given by Eq. (14.83) to Eq. (14.85) and the prescriptions described above, provides a particularly precise description of the solar  $\nu_e$  survival (and transitions) in the Sun - the results differ by a few percent from those obtained by solving numerically the relevant system of evolution equations using the electron number density distribution in the Sun provided by the standard solar models - if one uses as input in the calculations a “running” value of the scale-height  $r_0$  [173], *i.e.*, if for each given values of  $E/\Delta m_{21}^2$  and  $\theta_{12}$  one finds the resonance density  $N_e^{res} = N_e^{res}(E/\Delta m_{21}^2, \theta_{12})$ , calculates the scale-height parameter  $r_0 = N_e(r)/(dN_e(r)/dr)$  at the point in the Sun where  $N_e \cos^2 \theta_{13} = N_e^{res}(E/\Delta m_{21}^2, \theta_{12})$  employing the solar electron number density distribution  $N_e = N_e(r)$  given by the standard solar models [169],  $r$  being the distance from the center of the Sun.

#### 14.8.2.3. The day-night asymmetry:

When the solar neutrinos reaching a detector travel through the Earth at night, a partial regeneration of the flux of the solar  $\nu_e$  is possible due to the inverse Earth matter-enhanced process [183,184]  $\nu_{\mu(\tau)} \rightarrow \nu_e$ . This can lead to a difference between the solar neutrino induced charged current day and night event rates in the detector,  $R_D$  and  $R_N$ , *i.e.*, to a non-zero day-night asymmetry  $A_{D-N} = 2(R_D - R_N)/(R_D + R_N)$ . An observation of  $A_{D-N} \neq 0$  will be an unambiguous proof of the presence of Earth matter effects in the transitions of solar neutrinos taking place when the neutrinos traverse the Earth: in the absence of the effects of the Earth matter we have  $A_{D-N} = 0$ .

In the case of two-neutrino mixing, *i.e.*, neglecting the effects of the non-zero  $\sin \theta_{13}$ , the probability that an electron neutrino produced in the Sun will not be converted into  $\nu_{\mu(\tau)}$  when it propagates in the Sun and traverses the Earth on the way to the detector is given by the following simple expression [183]:

$$P_{SE}^{2\nu}(\nu_e \rightarrow \nu_e) = \bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e) + (1 - 2 \bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e)) \frac{P_{e2} - \sin^2 \theta_{12}}{\cos 2\theta_{12}}, \quad (14.90)$$

where  $\bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e)$  is the average probability of solar  $\nu_e$  survival in the Sun given in Eq. (14.84) and Eq. (14.85) (with  $\theta = \theta_{12}$  and  $\Delta m^2 = \Delta m_{21}^2 > 0$ ) and  $P_{e2} = |A(\nu_2 \rightarrow \nu_e)|^2$  is the probability of the  $\nu_2 \rightarrow \nu_e$  transition after the  $\nu_e$  have left the Sun, *i.e.*, of the  $\nu_2 \rightarrow \nu_e$  transition in the Earth. For solar neutrinos crossing only the Earth mantle along a trajectory with nadir angle  $\theta_n$ , the amplitude  $A(\nu_2 \rightarrow \nu_e)$  in the constant density approximation, has the form:

$$A(\nu_2 \rightarrow \nu_e) = \sin \theta_{12} + (e^{-i\varphi^{man}} - 1) \cos(\theta_{12} - \theta_{12}^{man}) \sin \theta_{12}^{man}, \quad (14.91)$$

where  $\varphi^{man} = (E_2^m - E_1^m)_{man} 2L^{man}$ ,  $(E_2^m - E_1^m)_{man}$  being the relevant difference of the energies of the two matter-eigenstate neutrinos in the Earth mantle and  $\theta_{12}^{man}$  is the mixing angle in the mantle which coincides in vacuum with  $\theta_{12}$ . The quantities  $(E_2^m - E_1^m)_{man}$  and  $\theta_{12}^{man}$  can be obtained from Eq. (14.59) and Eq. (14.57) by setting  $\theta = \theta_{12}$ ,  $\Delta m^2 = \Delta m_{21}^2$ ,  $N_e^{res} = N_{e21}^{res} = \Delta m_{21}^2 \cos 2\theta_{12}/(2E\sqrt{2}G_F)$  and  $N_e = \tilde{N}_e^{man}(\theta_n)$ . The two layer constant density approximation expressions for  $A(\nu_2 \rightarrow \nu_e)$  and  $P_{e2}$

for solar neutrinos crossing the Earth core at night were derived and can be found in Ref. 154.

During the day, when the neutrinos do not cross the Earth,  $P_{e2} = \sin^2 \theta_{12}$  and we have  $P_{SE}^{2\nu}(\nu_e \rightarrow \nu_e) = \bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e)$ . For Earth crossing neutrinos at night  $P_{e2} \neq \sin^2 \theta_{12}$  due to the Earth matter effect and  $P_{SE}^{2\nu}(\nu_e \rightarrow \nu_e) \neq \bar{P}_{\odot}^{2\nu}(\nu_e \rightarrow \nu_e)$ .

Detailed calculations of the day-night asymmetry  $A_{D-N} \neq 0$  for the solar neutrino detectors Super-Kamiokande, SNO and BOREXINO have been performed, *e.g.*, in Refs. [185]. In Refs. [186] the effects of a  $\theta_{13} \neq 0$  on the predictions for the asymmetry  $A_{D-N}$  were taken into account. The results of these calculations showed that for the experimentally determined current values of  $\Delta m_{21}^2$  and  $\theta_{12}$ , the predicted values of the asymmetry  $A_{D-N}$  for the SNO and BOREXINO experiments are below the sensitivity of these experiments. For the Super-Kamiokande detector an asymmetry  $A_{D-N} \sim -3\%$  was predicted.

## 14.9. Neutrino oscillation experiments

### 14.9.1. Solar neutrino experiments :

So far, solar neutrinos have been observed by chlorine (Homestake) [6] and gallium (SAGE [8], GALLEX [9,10], and GNO [11]) radiochemical detectors, water Cherenkov detectors using light water (Kamiokande [187,7] and Super-Kamiokande [188–191]) and heavy water (SNO [13,14,192,193]), and liquid scintillation detectors (Borexino [194–198] and KamLAND [199,200]).

A pioneering solar neutrino experiment by R. Davis, Jr. and collaborators at Homestake using the  $^{37}\text{Cl} - ^{37}\text{Ar}$  method proposed by B. Pontecorvo [201] started in the late 1960s. This experiment exploited  $\nu_e$  absorption on  $^{37}\text{Cl}$  nuclei followed by the produced  $^{37}\text{Ar}$  decay through orbital  $e^-$  capture,

$$\nu_e + ^{37}\text{Cl} \rightarrow ^{37}\text{Ar} + e^- \quad (\text{threshold } 814 \text{ keV}). \quad (14.92)$$

Note that  $\nu_e$  absorption reactions on nuclei are CC reactions. The detector contained 615 tons of tetrachloroethylene,  $\text{C}_2\text{Cl}_4$ . The  $^{37}\text{Ar}$  atoms produced are radioactive, with a half life ( $\tau_{1/2}$ ) of 34.8 days. After an exposure of the detector for two to three times  $\tau_{1/2}$ , the reaction products were chemically extracted and introduced into a low-background proportional counter, where they were counted for a sufficiently long period to determine the exponentially decaying signal and a constant background. Solar-model calculations predict that the dominant contribution in the chlorine experiment came from  $^8\text{B}$  neutrinos, the second to the dominant from  $^7\text{Be}$  neutrinos, with *pep*,  $^{13}\text{N}$ , and  $^{15}\text{O}$  neutrinos also giving additional subdominant contributions.

Gallium experiments (GALLEX and GNO at Gran Sasso in Italy and SAGE at Baksan in Russia) utilized the reaction

$$\nu_e + ^{71}\text{Ga} \rightarrow ^{71}\text{Ge} + e^- \quad (\text{threshold } 233 \text{ keV}), \quad (14.93)$$

which is sensitive to the most abundant *pp* solar neutrinos. The solar-model calculations predict that more than 80% of the capture rate in gallium is due to low energy *pp* and  $^7\text{Be}$  solar neutrinos with the *pp* rate being about twice the  $^7\text{Be}$  rate. The  $^{71}\text{Ge}$  atoms decay through electron capture with a half life ( $\tau_{1/2}$ ) of 11.43 days. SAGE used approximately 50 tons of liquid gallium metal as a target. GALLEX used 101 tons of  $\text{GaCl}_3$ , containing 30.3 tons of gallium. Both experiments used natural gallium, containing 39.9% of  $^{71}\text{Ga}$  isotope. SAGE started measurement from December, 1989. GALLEX experiment had been conducted between 1991 and 1997. Since April, 1998, a newly defined collaboration, GNO (Gallium Neutrino Observatory) continued the gallium experiment at Gran Sasso until April 2003.

Both GALLEX [202] and SAGE [203] tested their detectors using intense  $^{51}\text{Cr}$  radioactive sources with known activities. Low energy neutrinos relevant to test the gallium experiments ( $\sim 750 \text{ keV}$  and  $\sim 320 \text{ keV}$  neutrinos) are emitted from decays of  $^{51}\text{Cr}$ .

In 1987, the Kamiokande experiment at Kamioka in Japan succeeded in real-time solar neutrino observation, utilizing  $\nu_e$  scattering,

$$\nu_x + e^- \rightarrow \nu_x + e^-, \quad (14.94)$$



in a 3,000-ton water-Cherenkov detector. This experiment took advantage of the directional correlation between the incoming neutrino and the recoil electron. This feature greatly helps the clear separation of the solar-neutrino signal from the background. The Kamiokande result gave the first direct evidence that neutrinos come from the direction of the Sun [187]. In 1996, the high-statistics Super-Kamiokande experiment [188–191] with a 50-kton water Cherenkov detector replaced the Kamiokande experiment. Due to the high thresholds (recoil-electron total energy of 7 MeV in Kamiokande and 4 MeV at present in Super-Kamiokande) the experiments observe pure  $^8\text{B}$  solar neutrinos. It should be noted that the reaction (Eq. (14.94)) is sensitive to all active neutrinos,  $x = e, \mu$ , and  $\tau$ . However, the sensitivity to  $\nu_\mu$  and  $\nu_\tau$  is smaller than the sensitivity to  $\nu_e$  since  $\sigma(\nu_{\mu,\tau}e) \approx 0.16\sigma(\nu_e e)$ .

In 1999, a new real time solar-neutrino experiment, SNO (Sudbury Neutrino Observatory), in Canada started observation. This experiment used 1000 tons of ultra-pure heavy water ( $\text{D}_2\text{O}$ ) contained in a spherical acrylic vessel, surrounded by an ultra-pure  $\text{H}_2\text{O}$  shield. SNO measured  $^8\text{B}$  solar neutrinos via the CC and NC reactions

$$\nu_e + d \rightarrow e^- + p + p \quad (\text{CC}), \quad (14.95)$$

and

$$\nu_x + d \rightarrow \nu_x + p + n \quad (\text{NC}), \quad (14.96)$$

as well as  $\nu e$  scattering, (Eq. (14.94)). The CC reaction, (Eq. (14.95)), is sensitive only to  $\nu_e$ , while the NC reaction, (Eq. (14.96)), is sensitive to all active neutrinos. This is a key feature to solve the solar neutrino problem. If it is caused by flavour transitions such as neutrino oscillations, the solar neutrino fluxes measured by CC and NC reactions would show a significant difference.

The  $Q$ -value of the CC reaction is  $-1.4$  MeV and the  $e^-$  energy is strongly correlated with the  $\nu_e$  energy. Thus, the CC reaction provides an accurate measure of the shape of the  $^8\text{B}$  neutrino spectrum. The contributions from the CC reaction and  $\nu e$  scattering can be distinguished by using different  $\cos \theta$  distributions, where  $\theta$  is the angle of the  $e^-$  momentum with respect to the Sun-Earth axis. While the  $\nu e$  scattering events have a strong forward peak, CC events have an approximate angular distribution of  $(1 - 1/3 \cos \theta)$ .

The neutrino energy threshold of the NC reaction is 2.2 MeV. In the pure  $\text{D}_2\text{O}$  [13,14], the signal of the NC reaction was neutron capture in deuterium, producing a 6.25-MeV  $\gamma$ -ray. In this case, the capture efficiency was low and the deposited energy was close to the detection threshold of 5 MeV. In order to enhance both the capture efficiency and the total  $\gamma$ -ray energy (8.6 MeV), 2 tons of NaCl was added to the heavy water in the second phase of the experiment [192]. Subsequently NaCl was removed and an array of  $^3\text{He}$  neutron counters were installed for the third phase measurement [193]. These neutron counters provided independent NC measurement with different systematics from that of the second phase, and thus strengthened the reliability of the NC measurement. The SNO experiment completed data acquisition in 2006.

Another real time solar neutrino experiment, Borexino at Gran Sasso, started solar neutrino observation in 2007. This experiment measures solar neutrinos via  $\nu e$  scattering in 300 tons of ultra-pure liquid scintillator. With a detection threshold as low as 250 keV, the flux of monochromatic 0.862 MeV  $^7\text{Be}$  solar neutrinos has been directly observed for the first time [194]. Further, Borexino measured the fluxes of monochromatic 1.44 MeV  $\text{pep}$  solar neutrinos [195] and  $\text{pp}$  solar neutrinos [196], both for the first time. Measurements of these low energy solar neutrinos are important not only to test the SSM further, but also to study the MSW effect over the energy region spanning from sub-MeV to 10 MeV.

KamLAND is a 1-kton ultra-pure liquid scintillator detector located at the old Kamiokande's site. Recently this experiment also measured the  $^7\text{Be}$  solar neutrino flux [199]. As KamLAND is a multi-purpose experiment with one of the primary goals to be a long-baseline neutrino oscillation studies using electron antineutrinos emitted from nuclear power reactors, further description of the KamLAND experiment is given later in Section 14.9.4.

#### 14.9.2. Atmospheric neutrino oscillation experiments :

Almost all large underground detectors can observe atmospheric neutrinos. In the early history of neutrino oscillation studies using atmospheric neutrinos, water Cherenkov detectors for Kamiokande [204,205] and IMB [206] (experiment in the US) and iron tracking calorimeters for the Frejus experiment [207] in France and the Soudan 2 experiment [208] at the Soudan mine in the US, measured atmospheric neutrinos, in particular, the  $\Phi(\nu_\mu + \bar{\nu}_\mu)/\Phi(\nu_e + \bar{\nu}_e)$  ratio. The main purpose of all these experiments was search for nucleon decay, and atmospheric neutrinos were backgrounds for the main purpose. Following these initial experiments, Super-Kamiokande discovered the atmospheric neutrino oscillation [17], and a multi-purpose detector MACRO [209] at Gran Sasso obtained results consistent with neutrino oscillation. Later, the far detector of the MINOS long baseline neutrino oscillation experiment also measured atmospheric neutrinos [210], and obtained results consistent with atmospheric neutrino oscillation. This detector is a 5.4 kton iron-scintillator tracking calorimeter with toroidal magnetic field.

Atmospheric neutrino oscillations have also been observed by the neutrino telescopes for high-energy neutrino astronomy (TeV  $\sim$  PeV) using Cherenkov technique, ANTARES and IceCube-DeepCore, based on the measurement of  $\nu_\mu$  charged-current events having an upward-going muon track to avoid contamination from atmospheric muon background. ANTARES [211] is an open water detector deployed deep under the Mediterranean Sea (depth  $\sim$  2500 m) 40 km off-shore from Toulon, France, while IceCube [212] is a detector deployed in the ice at the South Pole at the depth from 1450 m to 2450 m. Though both experiments are optimized to high-energy neutrino interactions in the TeV range, they need to measure muon neutrinos with energies as low as  $\sim$  20 GeV in order to be sensitive to atmospheric neutrino oscillations. ANTARES could reconstruct upward-going muons from  $\nu_\mu$  interactions down to 20 GeV [213], while IceCube used the low-energy sub-detector DeepCore, a region of denser IceCube instrumentation, to lower the muon neutrino energy threshold down to 10 GeV [214] (also see Ref. 215 with 20 GeV threshold).

All these detectors, with the exception of the MINOS far detector, cannot measure the charge of the final-state leptons, and, therefore, neutrino and antineutrino induced events cannot be discriminated; the MINOS far detector can measure the charge of the muon track, and, therefore, identify  $\nu_\mu$  and  $\bar{\nu}_\mu$  charged-current events. However, all these detectors can identify the final-state leptons to be  $\mu$ -like or  $e$ -like. Taking Super-Kamiokande as an example, neutrino events having their vertex in the 22.5 kton fiducial volume are classified into fully contained (FC) events and partially contained (PC) events. The FC events are required to have no activity in the anti-counter. Single-ring events have only one charged lepton which radiates Cherenkov light in the final state, and particle identification is particularly clean for single-ring FC events. A ring produced by an  $e$ -like ( $e^\pm, \gamma$ ) particle exhibits a more diffuse pattern than that produced by a  $\mu$ -like ( $\mu^\pm, \pi^\pm$ ) particle, since an  $e$ -like particle produces an electromagnetic shower and low-energy electrons suffer considerable multiple Coulomb scattering in water. All the PC events are assumed to be  $\mu$ -like since the PC events comprise a 98% pure charged-current  $\nu_\mu$  sample.

In the near future, Super-Kamiokande and the MINOS (now MINOS+ [216]) far detector will continue atmospheric neutrino measurements. In addition, currently several large underground detectors are proposed for construction (liquid argon detectors with a total mass of 10 - 40 kton as the far detector of the DUNE experiment [84] in the US, and a 1 Mton water Cherenkov detector, Hyper-Kamiokande [165], as the far detector of the T2HK experiment [85] in Japan) or approved (a 50 kton magnetized iron tracking calorimeter, ICAL at the INO (India-based Neutrino Observatory) [219] in the southern Indian state of Tamil Nadu).

As high-statistics atmospheric neutrino observations in the energy region of a few to  $\sim$  10 GeV are considered to be promising for the determination of the neutrino mass ordering, see Section 14.8, there are two proposed densely-instrumented neutrino telescopes PINGU (Precision IceCube Next-Generation Upgrade) [220] and ORCA (Oscillation Research with Cosmics in the Abyss) [221], both having a multi-megaton total mass. PINGU will be deployed inside

the DeepCore, and will be sensitive to  $> 1$  GeV. ORCA is proposed as part of the second phase of the KM3NeT [222], a network of neutrino telescopes deep under the Mediterranean Sea. ORCA will have sensitivity down to a few GeV, its site being Toulon, France.

#### 14.9.3. Accelerator neutrino oscillation experiments :

For earlier accelerator neutrino oscillation experiments before the discovery of the atmospheric neutrino oscillation, see, *e.g.*, Ref. 223. The  $\Delta m^2 \geq 2 \times 10^{-3} \text{ eV}^2$  region can be explored by accelerator-based long-baseline experiments with typically  $E \sim 1$  GeV and  $L \sim$  several hundred km. With a fixed baseline distance and a narrower, well understood neutrino energy spectrum, the value of  $|\Delta m_{31(32)}^2|$  and, with higher statistics, also the relevant neutrino mixing angle, are potentially better constrained in accelerator experiments than from atmospheric neutrino observations. With  $\nu_\mu \rightarrow \nu_e$  appearance measurements, accelerator long-baseline experiments can measure  $\theta_{13}$  within an uncertainty related, in particular, to the CP-violating phase  $\delta$ . Using precise results on  $\theta_{13}$  from reactor experiments, they can potentially determine or constrain  $\delta$  and the neutrino mass ordering, depending on the experimental conditions such as baseline distance. K2K [19], MINOS [20,21] and MINOS+ [216], OPERA [141,142], ICARUS [224], T2K [23,22], and NO $\nu$ A [55,56] are completed or currently running experiments, and DUNE [84] and T2HK [85] are proposed future experiments.

The K2K (KEK-to-Kamioka) long-baseline neutrino oscillation experiment [19] is the first accelerator-based experiment with a neutrino path length extending hundreds of kilometers. A horn-focused wide-band muon neutrino beam having an average  $L/E_\nu \sim 200$  ( $L = 250$  km,  $\langle E_\nu \rangle \sim 1.3$  GeV), was produced by 12-GeV protons from the KEK-PS and directed to the Super-Kamiokande detector. A near detector was located 300 m downstream of the production target. K2K experiment started data-taking in 1999 and was completed in 2004.

MINOS [20,21] is the second long-baseline neutrino oscillation experiment with near and far detectors. Neutrinos are produced by the NuMI (Neutrinos at the Main Injector) facility using 120 GeV protons from the Fermilab Main Injector. The far detector is a 5.4 kton (total mass) iron-scintillator tracking calorimeter with toroidal magnetic field, located underground in the Soudan mine. The baseline distance is 735 km. The near detector, located 1.04 km downstream of the production target, is also an iron-scintillator tracking calorimeter with toroidal magnetic field, with a total mass of 0.98 kton. The NuMI neutrino beam is a horn-focused wide-band beam. Its energy spectrum can be varied by moving the target position relative to the first horn and changing the horn current. MINOS started the neutrino-beam run in 2005 and was completed in 2012. Almost all the MINOS data were taken with the low-energy beam spectrum which peaked at 3 GeV. Part of the MINOS data were taken with the  $\bar{\nu}_\mu$ -enhanced beam by inverting the current in magnetic horns. In September, 2013, the MINOS+ experiment [216] started with the same near and far detectors as the MINOS experiment, but with the medium-energy beam spectrum which peaks at 7 GeV. At zero degree, the NuMI medium-energy beam has much higher intensity than the NuMI low-energy beam.

The T2K experiment [22,23] is the first off-axis long-baseline neutrino oscillation experiment. The baseline distance is 295 km between the J-PARC in Tokai, Japan and Super-Kamiokande. A narrow-band  $\nu_\mu$  beam with a peak energy of 0.6 GeV, produced by 30 GeV protons from the J-PARC Main Ring, is directed  $2.5^\circ$  off-axis to SK. With this configuration, the  $\nu_\mu$  beam is tuned to the first oscillation minimum of the  $\nu_\mu$  survival probability. T2K started the first physics run in 2010.

The NO $\nu$ A experiment [55,56] is an off-axis long-baseline neutrino oscillation experiment using the the NuMI medium-energy beam. Its detectors are positioned 14 mrad off-axis. With this configuration, the neutrino beam has a narrow spectrum which peaks at 2 GeV. The 14 kton far detector is located on the surface at Ash River, Minnesota, 810 km from the production target. The 0.3 kton near detector is located underground at Fermilab, approximately 1km from the target. Both detectors are fine-grained tracking calorimeters consisting of

arrays of PVC cells filled with liquid scintillator. NO $\nu$ A started full operation in October, 2014.

Although the atmospheric neutrino oscillations and accelerator long-baseline  $\nu_\mu$  disappearance data are fully consistent with the dominance of  $\nu_\mu \rightarrow \nu_\tau$  oscillations for  $\nu_\mu$  at GeV energies,  $\nu_\tau$  appearance in the muon neutrino beam has to be demonstrated. As the  $\tau$  production threshold is  $E_\nu \sim 3.5$  GeV, a high-energy neutrino beam is needed for this purpose. The only experiment of this kind is OPERA [141,142] with a muon neutrino source at CERN and a detector at Gran Sasso with the baseline distance of 730 km. OPERA does not have a near detector. The CNGS (CERN Neutrinos to Gran Sasso) neutrino beam with  $\langle E_\nu \rangle = 17$  GeV is produced by high-energy protons from the CERN SPS. OPERA received the CNGS neutrino beam between 2008 and 2012. The detector is a combination of the “Emulsion Cloud Chamber” and magnetized spectrometer, having a target mass of 1,290 tons. At Gran Sasso, another neutrino experiment, ICARUS [224], with a 600-ton liquid argon detector, was located and received the CNGS neutrino beam from 2010 to 2012. The ICARUS detector will be transported to Fermilab by 2017, and will be used in a short baseline experiment.

DUNE (Deep Underground Neutrino Experiment) [84] is a projected future experiment with a 1,300 km baseline. A  $10 \sim 34$  kton liquid-argon far detector will be located deep underground at the Sanford Lab in South Dakota, the U.S. A fine-grained near neutrino detector will be installed at Fermilab. Based on the existing NuMI beamline and a MW class proton source, a wide-band, high-intensity  $\nu_\mu$  beam with a peak flux at 2.5 GeV is considered for this experiment. T2HK [85] is another future long baseline experiment from J-PARC to the 1 Mton water Cherenkov detector, Hyper-Kamiokande [165], which is at the proposal stage, at Kamioka. An upgrade of the J-PARC Main Ring to achieve a MW-class beam power is also proposed.

In the context of possible hints for the existence of sterile neutrinos at the eV scale, short-baseline accelerator neutrino oscillation experiments have been drawing attention. LSND [225], Karmen 2 (and Karmen 1) [226], and MiniBooNE [227,228] are completed experiments. New short-baseline experiments, MicroBooNE, SBND, and ICARUS are in preparation at Fermilab [229]. The detectors of all these new experiments use liquid-argon TPC technology.

The LSND (Liquid Scintillation Neutrino Detector) experiment [225] used the LANSE (Los Alamos Neutron Science Center, formerly known as the LAMPF) 800 MeV proton linac as a neutrino source. At this energy, kaon production is negligible. Most of the produced positive pions stop in the massive target and decay at rest, with decay muons also stop in the target and decay. Most of the produced negative pions also stop in the target and are absorbed by the target nuclei. Therefore, this neutrino source emits  $\nu_\mu$ ,  $\bar{\nu}_\mu$ , and  $\nu_e$ , with very small contamination of  $\bar{\nu}_e$  which comes from  $\pi^-$  decay in flight followed by the  $\mu^-$  decay at rest. Because of this small  $\bar{\nu}_e$  component in the neutrino flux, LSND made a sensitive search for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance with 167 tons of diluted liquid scintillator in a tank located about 30 m from the neutrino source, using the reaction  $\bar{\nu}_e + p \rightarrow e^+ + n$ . Also, LSND studied  $\nu_\mu \rightarrow \nu_e$  appearance above the Michel electron endpoint energy using the reaction  $\nu_e + C \rightarrow e^- + N$ , as the  $\nu_e$  flux from  $\mu^+$  decay in flight is suppressed due to the long muon lifetime and that from  $\pi^+$  decay in flight is suppressed by the small  $\pi^+ \rightarrow e^+ + \nu_e$  branching ratio. The Karmen 2 experiment [226] used the 800 MeV proton synchrotron at the neutron spallation facility of the Rutherford Appleton Laboratory, also to produce low-energy  $\bar{\nu}_\mu$  flux from  $\mu^+$  decay at rest. The Karmen 2 detector is a segmented liquid scintillator calorimeter located at a distance of 17.7 m from the neutrino source. MiniBooNE used a conventional horn-focused neutrino beam produced by 8 GeV protons from the Fermilab booster synchrotron. MiniBooNE investigated both  $\nu_e$  [227] and  $\bar{\nu}_e$  [228] appearance in  $\nu_\mu$  and  $\bar{\nu}_\mu$  beams, respectively, with a detector containing 806 tons of mineral oil and located 541 m downstream of the production target.

#### 14.9.4. Reactor neutrino oscillation experiments :

As the relevant energy range of the events detected in the reactor experiment is  $< 10$  MeV, detection of a prompt positron signal and delayed neutron signal in coincidence is important to identify the inverse  $\beta$ -decay, rejecting natural backgrounds. For detecting neutrons effectively, gadolinium-loaded liquid scintillator is widely used. While neutron capture on a hydrogen produces a 2.2 MeV  $\gamma$ , neutron capture on Gd produces multiple  $\gamma$ , each having average energy of  $\sim 2$  MeV, giving a  $\sim 8$  MeV signal in total.

For short baseline reactor  $\bar{\nu}_e$  neutrino oscillation experiments in 1980s or earlier, see, *e.g.*, Refs. [230,46]. Reactor  $\bar{\nu}_e$  disappearance experiments with  $L \sim 1$  km,  $\langle E \rangle \sim 3$  MeV are sensitive to  $E/L \sim 3 \times 10^{-3} \text{ eV}^2 \sim |\Delta m_{31(32)}^2|$ . At this baseline distance, the reactor  $\bar{\nu}_e$  oscillations driven by  $\Delta m_{21}^2$  are negligible. Therefore, as can be seen from Eq. (14.44) and Eq. (14.46),  $\theta_{13}$  can be directly measured. An experiment at the Chooz Nuclear Power Station in France [51] was the first experiment of this kind. The detector was located in an underground laboratory with 300 mwe (meter water equivalent) rock overburden, at about 1 km from the neutrino source. It consisted of a central 5-ton target filled with 0.09% Gd-loaded liquid scintillator, surrounded by an intermediate 17-ton and outer 90-ton regions filled with Gd-free liquid scintillator. Another experiment at the Palo Verde Nuclear Generating Station in Arizona, United States, also searched for  $\bar{\nu}_e$  disappearance using 11.34 tons of Gd-loaded liquid scintillator located at a shallow (32 mwe) underground site, about 800 m from the neutrino source [231].

Although these two experiments in the late 1990s had found no evidence for  $\bar{\nu}_e$  disappearance, after establishment of atmospheric and solar neutrino oscillations the importance of the reactor neutrino oscillation experiment to measure  $\theta_{13}$  was widely recognized, and this led to the realization of the three new reactor experiments, Double Chooz in France, RENO in Korea, and Daya Bay in China. The Double Chooz experiment [30,33] measures electron antineutrinos from two 4.25 GW<sub>th</sub> reactors with a far detector at an average distance of 1050 m from the two reactor cores. The Daya Bay experiment [31,35] measures  $\bar{\nu}_e$ s from the Daya Bay nuclear power complex (six 2.9 GW<sub>th</sub> reactors), initially with six functionally identical detectors deployed in two near (470 m and 576 m of flux-weighted baselines) and one far (1648 m) underground halls. The first Daya Bay result [31] was obtained with this detector configuration. Later, two detectors were further installed, one in one of the near detector hall and the other in the far detector hall. The RENO experiment [32] measures electron antineutrinos from four 2.8 GW<sub>th</sub> and two 2.66 GW<sub>th</sub> reactors at Yonggwang Nuclear Power Plant with two identical detectors located at 294 m and 1383 m from the reactor array center (or flux-weighted baseline distance of 408.56 m and 1443.99 m, respectively). Antineutrino detectors of these experiments have similar structures. They consist of three layers and an optically independent outer veto detector. The innermost layer of the antineutrino detector is filled with Gd-loaded liquid scintillator, which is surrounded by a “ $\gamma$ -catcher” layer filled with Gd-free liquid scintillator, and outside the  $\gamma$ -catcher is a buffer layer filled with mineral oil. An outer veto detector is filled with purified water (Daya Bay and RENO) or liquid scintillator (Double Chooz). In addition, the Double Chooz near detector tank is shielded by a 1 m thick water buffer. RENO and Daya Bay started measurements with both the near and far detectors from the beginning. However, the Double Chooz near detector has been completed at the end of 2014. All these experiments published their first results on reactor  $\bar{\nu}_e$  disappearance in 2012.

For longer baseline distance of  $L \sim$  a few hundred km, a reactor neutrino oscillation experiment is sensitive to  $\Delta m^2$  down to  $\sim 10^{-5} \text{ eV}^2$ . Therefore, such an experiment can test the LMA (Large Mixing Angle) solution of the solar neutrino problem, assuming CPT invariance. However, a higher  $\bar{\nu}_e$  flux and a larger target mass are needed compared to short-baseline reactor experiments to obtain statistically significant event rate. So far, KamLAND is the only experiment of this kind. It is located at the old Kamiokande’s site in Japan. Its neutrino target is 1-kton ultra-pure liquid scintillator contained in a transparent balloon, which is hold inside a spherical tank with buffer oil filled between the balloon and the tank. The tank

is surrounded by an outer water Cherenkov detector. Before the Great East Japan Earthquake in March 2011, many nuclear reactors were operating in Japan, and more than 79% of the  $\bar{\nu}_e$  flux at KamLAND was coming from 26 reactors between 138 – 214 km away, with a flux-weighted average distance of  $\sim 180$  km.

In future, medium baseline ( $\sim 50$  km) reactor neutrino oscillation experiments with neutrino target mass of  $\sim 20$  kton and with a very good energy resolution of  $3\%/\sqrt{E_\nu(\text{MeV})}$ , not reached in any previous experiment with liquid scintillator, are aiming, in particular, to determine the type of spectrum the neutrino masses obey, *i.e.*, the neutrino mass ordering (see Section 14.2). These experiments have additional rich physics program. The Jiangmen Underground Neutrino Observatory (JUNO) [232] at Kaiping, Jiangmen in Southern China will be located at 53 km from both of the planned Yangjiang and Taishan nuclear power plants. The neutrino target of this experiment will be 20 kton liquid scintillator. JUNO is a funded project, and its construction started in January, 2015. A similar experiment, RENO50 [127], is also proposed in Korea.

### 14.10. Results of solar neutrino experiments and KamLAND

In 1967, analyzing the possible effects of neutrino oscillations on the solar neutrino flux measurements, B. Pontecorvo predicted the solar neutrino “deficit” in experiments detecting solar neutrinos via a CC reaction [1] before the first solar neutrino data were available. The solar-neutrino problem, *i.e.*, the problem of understanding the origin of the observed deficit of solar neutrinos, remained unsolved for more than 30 years since the late 1960s, but solar neutrino experiments have achieved remarkable progress since the beginning of the new century, and the solar-neutrino problem has been understood as due to neutrino flavour conversion.

#### 14.10.1. Measurements of $\Delta m_{21}^2$ and $\theta_{12}$ :

From the very beginning of the solar-neutrino observation by the Homestake chlorine experiment [233] in the late 1960s, it was recognized that the observed flux was significantly smaller than the SSM prediction. The subsequent radiochemical solar neutrino experiments using  $^{71}\text{Ga}$ , SAGE [234] and GALLEX [9] also reported smaller solar neutrino fluxes than the SSM predictions in the early 1990s. final results of these experiments [235] are compared with the SSM predictions in Table 14.5. Experiments with water Cherenkov detectors, Kamiokande and Super-Kamiokande, observed almost pure  $^8\text{B}$  solar neutrinos through  $\nu e$  elastic scattering, and they also reported a clear deficit of  $^8\text{B}$  solar neutrino flux.

**Table 14.5:** Results from radiochemical solar-neutrino experiments. The predictions of the standard solar model BPS08(GS) are also shown. The first and the second errors in the experimental results are the statistical and systematic errors, respectively. SNU (Solar Neutrino Unit) is defined as  $10^{-36}$  neutrino captures per atom per second.

	$^{37}\text{Cl} \rightarrow ^{37}\text{Ar}$ (SNU)	$^{71}\text{Ga} \rightarrow ^{71}\text{Ge}$ (SNU)
Homestake [6]	$2.56 \pm 0.16 \pm 0.16$	–
GALLEX [10]	–	$77.5 \pm 6.2^{+4.3}_{-4.7}$
GNO [11]	–	$62.9^{+5.5}_{-5.3} \pm 2.5$
GNO+GALLEX [11]	–	$69.3 \pm 4.1 \pm 3.6$
SAGE [8]	–	$65.4^{+3.1+2.6}_{-3.0-2.8}$
SSM [BPS08(GS)] [100]	$8.46^{+0.87}_{-0.88}$	$127.9^{+8.1}_{-8.2}$

In 2001, the initial SNO CC result combined with the Super-Kamiokande’s high-statistics  $\nu e$  elastic scattering result [236] provided direct evidence for flavour conversion of solar neutrinos [13]. Later, SNO’s NC measurements further strengthened this conclusion [14,192,193]. From the salt-phase measurement [192], the fluxes

measured with CC, ES, and NC events were obtained as

$$\phi_{\text{SNO}}^{\text{CC}} = (1.68 \pm 0.06^{+0.08}_{-0.09}) \times 10^6 \text{cm}^{-2} \text{s}^{-1}, \quad (14.97)$$

$$\phi_{\text{SNO}}^{\text{ES}} = (2.35 \pm 0.22 \pm 0.15) \times 10^6 \text{cm}^{-2} \text{s}^{-1}, \quad (14.98)$$

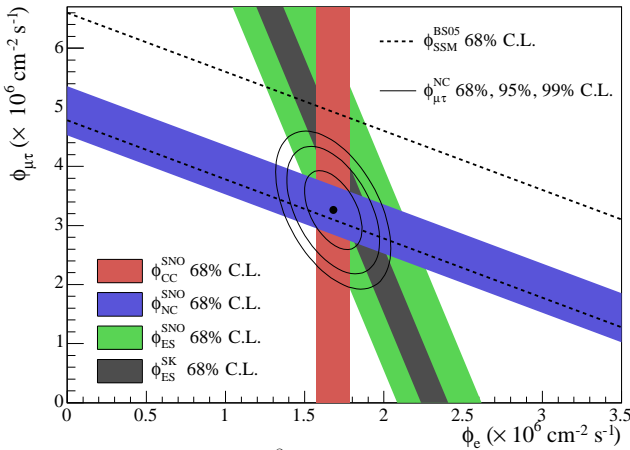
$$\phi_{\text{SNO}}^{\text{NC}} = (4.94 \pm 0.21^{+0.38}_{-0.34}) \times 10^6 \text{cm}^{-2} \text{s}^{-1}, \quad (14.99)$$

where the first errors are statistical and the second errors are systematic. In the case of  $\nu_e \rightarrow \nu_{\mu,\tau}$  transitions, Eq. (14.99) is a mixing-independent result and therefore tests solar models. It shows good agreement with the  $^8\text{B}$  solar-neutrino flux predicted by the solar model [98]. Fig. 14.8 shows the salt phase result on the  $\nu_{\mu} + \nu_{\tau}$  flux  $\phi(\nu_{\mu,\tau})$  versus the flux of electron neutrinos  $\phi(\nu_e)$  with the 68%, 95%, and 99% joint probability contours. The flux of non- $\nu_e$  active neutrinos,  $\phi(\nu_{\mu,\tau})$ , can be deduced from these results. It is

$$\phi(\nu_{\mu,\tau}) = (3.26 \pm 0.25^{+0.40}_{-0.35}) \times 10^6 \text{cm}^{-2} \text{s}^{-1}. \quad (14.100)$$

The non-zero  $\phi(\nu_{\mu,\tau})$  is strong evidence for neutrino flavor conversion. These results are consistent with those expected from the LMA (large mixing angle) solution of solar neutrino oscillation in matter [26,27] with  $\Delta m_{21}^2 \sim 7.5 \times 10^{-5} \text{eV}^2$  and  $\tan^2 \theta_{12} \sim 0.45$ . However, with the SNO data alone, the possibility of other solutions of solar neutrino oscillation in matter cannot be excluded with sufficient statistical significance.

The KamLAND experiment solved this problem and finally identified the LMA solution as the true solution of the solar neutrino problem. With the reactor  $\bar{\nu}_e$ 's energy spectrum ( $< 8 \text{ MeV}$ ) and a prompt-energy analysis threshold of  $2.6 \text{ MeV}$ , this experiment has a sensitive  $\Delta m^2$  range down to  $\sim 10^{-5} \text{eV}^2$ . Therefore, if the LMA solution is the real solution of the solar neutrino problem, KamLAND should observe reactor  $\bar{\nu}_e$  disappearance, assuming CPT invariance.



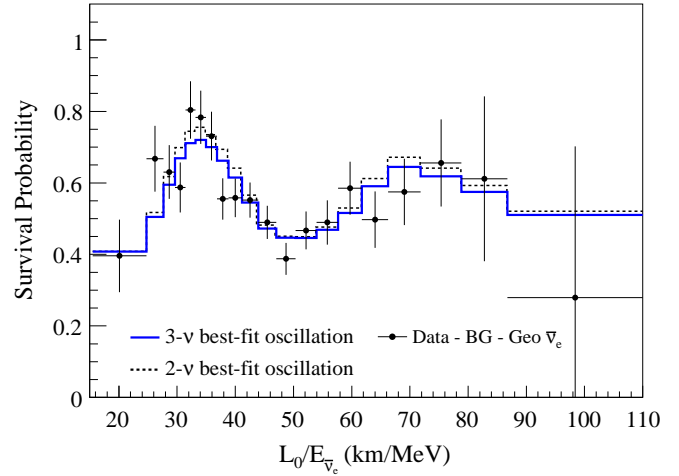
**Figure 14.8:** Fluxes of  $^8\text{B}$  solar neutrinos,  $\phi(\nu_e)$ , and  $\phi(\nu_{\mu,\tau})$ , deduced from the SNO's CC, ES, and NC results of the salt phase measurement [192]. The Super-Kamiokande ES flux is from Ref. 237. The BS05(OP) standard solar model prediction [98] is also shown. The bands represent the  $1\sigma$  error. The contours show the 68%, 95%, and 99% joint probability for  $\phi(\nu_e)$  and  $\phi(\nu_{\mu,\tau})$ . The figure is from Ref. 192.

The first KamLAND results [15] with 162 ton-yr exposure were reported in December 2002. The ratio of observed to expected (assuming no  $\bar{\nu}_e$  oscillations) number of events was

$$\frac{N_{\text{obs}} - N_{\text{BG}}}{N_{\text{NoOsc}}} = 0.611 \pm 0.085 \pm 0.041 \quad (14.101)$$

with obvious notation. This result showed clear evidence of an event deficit expected from neutrino oscillations. The 95% CL allowed regions are obtained from the oscillation analysis with the observed

event rates and positron spectrum shape. A combined global solar + KamLAND analysis showed that the LMA is a unique solution to the solar neutrino problem with  $> 5\sigma$  CL [238]. With increased statistics [16,239,240], KamLAND observed not only the distortion of the  $\bar{\nu}_e$  spectrum, but also for the first time the periodic dependence on the neutrino energy of the  $\bar{\nu}_e$  survival probability expected from neutrino oscillations (see Fig. 14.9).



**Figure 14.9:** The ratio of the background and geoneutrino-subtracted  $\bar{\nu}_e$  spectrum, observed in the KamLAND experiment, to the predicted one without oscillations (survival probability) as a function of  $L_0/E$ , where  $L_0=180 \text{ km}$ . The histograms show the expected distributions based on the best-fit parameter values from the two- and three-flavor neutrino oscillation analyses. The figure is from Ref. 240.

It should be noted that with accumulation of precise solar neutrino data, analyses using only solar neutrino data [241,242,198] have attained sufficiently high statistical significance ( $> 99.73\%$  or  $> 3\sigma$  CL) to show the LMA solution to be the real solution to the solar neutrino problem without resorting to the KamLAND data, namely, without assuming CPT invariance, though the allowed  $\Delta m_{21}^2$  range is better determined by the KamLAND data.

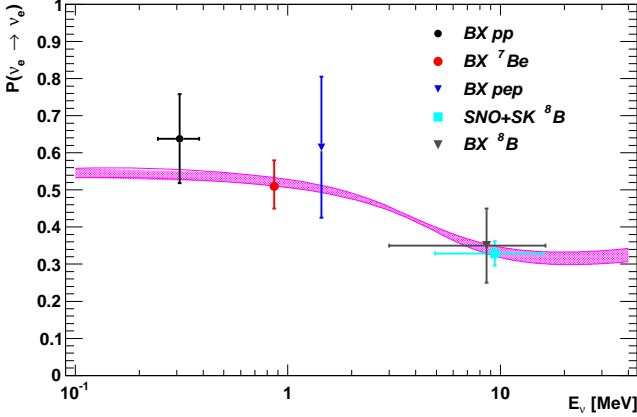
The values of  $\Delta m_{21}^2$  and  $\theta_{12}$  have been frequently updated by experimental groups or by phenomenological analysis groups, using the global solar neutrino data, or the KamLAND data alone, or the global solar + KamLAND data, or the global neutrino oscillation data. The latest global analysis results found in Ref. 60 are shown in Table 14.1.

Regarding the consistency between the KamLAND and solar neutrino experiments on the values of  $\Delta m_{21}^2$  and  $\theta_{12}$ , it has been noted that there is a  $\sim 2\sigma$  level tension between the best-fit value of  $\Delta m_{21}^2$  determined by the KamLAND collaboration and that obtained from analyses using global solar neutrino data [53]. The solar data prefer lower  $\Delta m_{21}^2$  value. The KamLAND and global solar best-fit values of  $\theta_{12}$  are consistent.

#### 14.10.2. Solar neutrino flux measurements and indications of matter effects :

So far, the pp, pep,  $^7\text{Be}$ ,  $^8\text{B}$  solar neutrino fluxes have been measured, and upper limits have been set for the hep and CNO solar neutrino fluxes, with various techniques. Chlorine (Homestake) and gallium (SAGE, GALLEX, and GNO) radiochemical experiments measured capture rates of solar neutrinos above threshold (see Table 14.5). Light-water Cherenkov detectors, Kamiokande [7] and Super-Kamiokande [189,191], measured the  $^8\text{B}$  neutrino flux and set an upper limit for the hep neutrino flux using  $\nu_e$  elastic scattering [189]. A heavy-water Cherenkov detector, SNO [242], also measured the  $^8\text{B}$  neutrino flux, but with three different reactions, NC, CC, and  $\nu e$  elastic scattering. Liquid scintillator detectors, Borexino and KamLAND, measured low-energy solar neutrinos using  $\nu e$  elastic

scattering. In particular, Borexino [198] successfully measured the pp [196], pep [195], and  ${}^7\text{Be}$  [194] solar neutrino fluxes and set an upper limit for the CNO solar neutrino flux [195]. KamLAND also measured the  ${}^7\text{Be}$  solar neutrino flux [199]. In addition, both Borexino [197] and KamLAND [200] measured the  ${}^8\text{B}$  neutrino flux. The measured fluxes or upper limits from all these experiments are listed in the Particle Listings of this RPP edition.



**Figure 14.10:** Electron neutrino survival probability as a function of neutrino energy according to the MSW-LMA model. The low-energy region ( $< 1$  MeV) of the curve is calculated for pp and  ${}^7\text{Be}$  neutrinos, and the high-energy region for  ${}^8\text{B}$  neutrinos, using the parameter values given in Ref. 52. The width of the curve reflects  $\pm 1\sigma$  uncertainties, determined by a Monte Carlo method sampling errors on parameters [52]. The points represent the Borexino pp,  ${}^7\text{Be}$ , pep, and  ${}^8\text{B}$  data and the SNO+SK  ${}^8\text{B}$  data. This figure is provided by A. Ianni in the name of the BOREXINO Collaboration.

Fig. 14.10 plots the survival probability of solar  $\nu_e$  as a function of neutrino energy. The data points are from the Borexino results except the SNO+SK  ${}^8\text{B}$  data. As explained in Section 14.8.2.2, matter effects on solar neutrino oscillation is expected to be given by the average two-neutrino vacuum oscillation probability,  $1 - \frac{1}{2}\sin^2 2\theta_{12}$  for survival of pp neutrinos. It is  $\sim 0.58$  for  $\sin^2 \theta_{12} = 0.297$ . For  ${}^8\text{B}$  neutrinos, transitions are adiabatic and the survival probability is given by  $\sin^4 \theta_{13} + \cos^4 \theta_{13} \sin^2 \theta_{12} \sim 0.28$  for  $\sin^2 \theta_{13} = 0.0214$  (for normal mass ordering) and  $\sin^2 \theta_{12} = 0.297$ . All the data shown in this plot are consistent with the theoretically calculated curve. This indicates that these solar neutrino results are consistent with the MSW-LMA solution of the solar neutrino problem.

In the nighttime, solar neutrino experiments observe neutrinos propagated through the Earth. Therefore, a non-zero day-night flux (or interaction rate) asymmetry implies the Earth matter effects on flavour oscillations of solar neutrinos. In particular, if the nighttime flux is higher than the daytime flux, it implies a  $\nu_e$  regeneration by the Earth matter effects (see Section 14.8.2.3). Previously, SNO [242] and Borexino [243] searched for day-night flux asymmetries of  ${}^8\text{B}$  and  ${}^7\text{Be}$  neutrinos, respectively, but they observed no statistically significant asymmetries. Recently, the Super-Kamiokande experiment has reported [244] a  $2.7\sigma$  indication of non-zero day-night asymmetry of  ${}^8\text{B}$  solar neutrinos,  $A_{DN} = 2(R_D - R_N)/(R_D + R_N) = -0.032 \pm 0.011 \pm 0.005$ , where  $R_D$  and  $R_N$  are the average day and average night  $\nu_e$  elastic-scattering rates of  ${}^8\text{B}$  solar neutrinos. This result is consistent with the  $\Delta m_{21}^2$  and  $\theta_{12}$  values in the LMA region.

#### 14.11. Measurements of $|\Delta m_{31(32)}^2|$ and $\theta_{23}$ , and related topics

The first compelling evidence for the neutrino oscillation was  $\nu_\mu$  disappearance observed by the Super-Kamiokande Collaboration in 1998 [17] in the measurement of atmospheric neutrinos produced by cosmic-ray interactions in the atmosphere. A striking feature of atmospheric neutrino oscillations was a surprisingly large mixing angle  $\theta_{23}$ . Whether mixing is maximal, *i.e.*,  $\theta_{23} = \pi/4$ , or, if not, in which octant  $\theta_{23}$  lies, is one of the questions drawing much interest in neutrino physics because the measurement of certain fundamental physical observables depends on the value of  $\sin^2 \theta_{23}$  (see, *e.g.*, Sections 14.2 and 14.8.1). The high precision measurement of  $\sin^2 \theta_{23}$  will provide also a test of a large class of theories of neutrino masses and mixing, based, in particular, on discrete symmetries (see, *e.g.*, the first two articles quoted in Ref. 87 and Ref. 245).

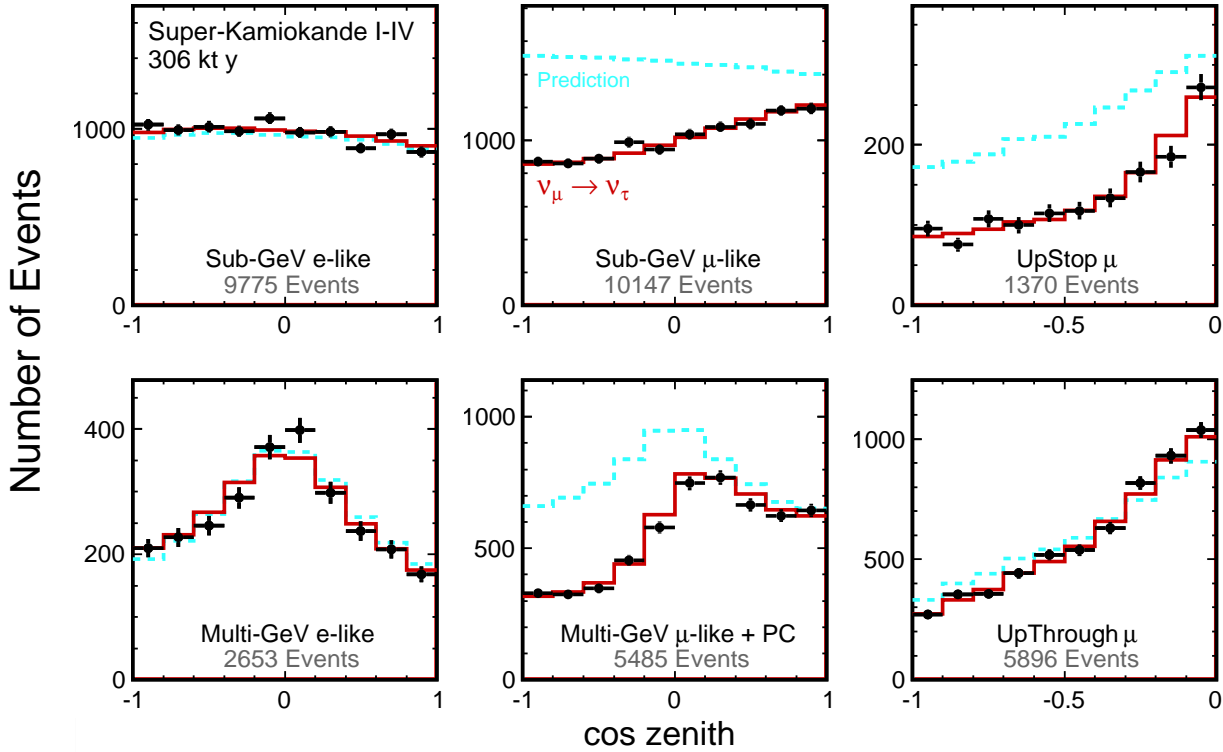
##### 14.11.1. $\nu_\mu$ disappearance :

Prior to the Super-Kamiokande's discovery of atmospheric neutrino oscillations, a deficit of atmospheric  $\nu_\mu + \bar{\nu}_\mu$  flux was indicated by the Kamiokande experiment [204]. Actually, Kamiokande reported the double ratio  $R(\nu_\mu/\nu_e) = (\text{Measured } \nu_\mu/\nu_e)/(\text{Expected } \nu_\mu/\nu_e) < 1$  to reduce systematic effects due to rather large flux uncertainties. The IMB [206] and Soudan 2 [208] experiments also observed  $R < 1$ , but the Frejus experiment [207] did not see such a tendency. Kamiokande further observed zenith-dependence of  $\nu_\mu + \bar{\nu}_\mu$  flux deficit [205]. However, all these results from early experiments did not have conclusive statistical significance.

In 1998, the Super-Kamiokande Collaboration reported a significant zenith-angle ( $\Theta$ ) dependent deficit of  $\mu$ -like events compared to the no-oscillation expectation [17]. For multi-GeV (visible energy  $> 1.33$  GeV) FC+PC muons, the asymmetry  $A$ , defined as  $A = (U - D)/(U + D)$ , where  $U$  is the number of upward-going events ( $-1 < \cos \Theta < -0.2$ ) and  $D$  is the number of downward-going events ( $0.2 < \cos \Theta < 1$ ), was observed to be  $A = -0.296 \pm 0.048 \pm 0.01$  which deviates from 0 by more than  $6\sigma$ . This asymmetry is expected to be  $\sim 0$  independent of the atmospheric neutrino flux model for neutrino energy  $> 1$  GeV. On the other hand, the zenith-angle distribution of the  $e$ -like events was consistent with the expectation in the absence of oscillations. Fig. 14.11 shows the recent compilation of zenith-angle distributions of  $e$ -like and  $\mu$ -like events from the Super-Kamiokande atmospheric observation. Events included in these plots are single-ring FC events subdivided into sub-GeV (visible energy  $< 1.33$  GeV) events and multi-GeV events. The zenith-angle distribution of the multi-GeV  $\mu$ -like events is shown combined with that of the PC events. The final-state leptons in these events have good directional correlation with the parent neutrinos. The dotted histograms show the Monte Carlo expectation for neutrino events. If the produced flux of atmospheric neutrinos of a given flavour remains unchanged at the detector, the data should have similar distributions to the expectation. However, the zenith-angle distribution of the  $\mu$ -like events shows a strong deviation from the expectation. On the other hand, the zenith-angle distribution of the  $e$ -like events is consistent with the expectation. This characteristic feature is interpreted that muon neutrinos coming from the opposite side of the Earth's atmosphere, having travelled  $\sim 10,000$  km, oscillate into other neutrinos and disappeared, while oscillations still do not take place for muon neutrinos coming from above the detector, having travelled from a few to a few tens km. These results are in good agreement with  $\nu_\mu \leftrightarrow \nu_\tau$  two-flavour neutrino oscillations, because there is no indication of electron neutrino appearance. The atmospheric neutrinos corresponding to the events shown in Fig. 14.11 have  $E = 1 \sim 10$  GeV. With  $L = 10000$  km, neutrino oscillations suggests  $\Delta m^2 \sim 10^{-3} \sim 10^{-4}$   $\text{eV}^2$ . A significant deficit of  $\mu$ -like events suggests a large mixing angle. Super-Kamiokande's initial results on the oscillation parameters for  $\nu_\mu \leftrightarrow \nu_\tau$  were  $5 \times 10^{-4} < \Delta m^2 < 6 \times 10^{-3}$   $\text{eV}^2$  and  $\sin^2 2\theta > 0.82$  at 90% CL [17].

Although the Super-Kamiokande's atmospheric neutrino data are consistent with  $\nu_\mu \leftrightarrow \nu_\tau$  oscillations, this interpretation will be strengthened if  $\nu_\tau$  appearance and characteristic sinusoidal behavior of the  $\nu_\mu$  survival probability as a function of  $L/E$  were observed. In fact, other exotic explanations such as neutrino decay [246] and quantum





**Fig. 14.11:** The zenith angle distributions for fully contained 1-ring  $e$ -like and  $\mu$ -like events with visible energy  $< 1.33$  GeV (sub-GeV) and  $> 1.33$  GeV (multi-GeV). For multi-GeV  $\mu$ -like events, a combined distribution with partially contained (PC) events is shown. The dotted histograms show the non-oscillated Monte Carlo events, and the solid histograms show the best-fit expectations for  $\nu_\mu \leftrightarrow \nu_\tau$  oscillations. (This figure is provided by the Super-Kamiokande Collab.)

decoherence [247] cannot be completely ruled out from the zenith-angle distributions alone. By selecting events with high  $L/E$  resolution, evidence for the dip in the  $L/E$  distribution was observed at the right place expected from the interpretation of the Super-Kamiokande's data in terms of  $\nu_\mu \leftrightarrow \nu_\tau$  oscillations [18], see Fig. 14.12. This dip cannot be explained by alternative hypotheses of neutrino decay and neutrino decoherence, and they are excluded at more than  $3\sigma$  in comparison with the neutrino oscillation interpretation. For  $\nu_\tau$  appearance, see Section 14.11.4.

The muon neutrino disappearance discovered by Super-Kamiokande has been confirmed subsequently by atmospheric neutrino experiments, MACRO [209] and Soudan 2 [248], long baseline accelerator experiments, K2K [19], MINOS [20,21], T2K [22,23], and NO $\nu$ A [55], and neutrino telescope experiments, ANTARES [211] and IceCube-DeepCore [212]. Fig. 14.13 shows 90% CL allowed regions in the  $\sin^2 2\theta_{23}$  -  $\Delta m_{32(31)}^2$  plane, for the case of normal mass ordering, reported by the T2K [249], MINOS [250], Super-Kamiokande [251], and IceCube-DeepCore [212] experiments. All these regions are derived from three-neutrino oscillation analyses.

#### 14.11.2. Octant of $\theta_{23}$ :

The two-flavour  $\nu_\mu$  survival probability in vacuum is degenerate with respect to the interchange  $\theta_{23} \leftrightarrow \pi/2 - \theta_{23}$ . In other words, the  $\nu_\mu$  disappearance is not sensitive to the octant of  $\theta_{23}$  in the leading order. To determine the octant of  $\theta_{23}$ , it is necessary to perform, *e.g.*, precise measurements of  $\nu_\mu$  disappearance and analyses in terms of three-neutrino oscillations, or combined analysis of  $\nu_\mu$  disappearance and  $\nu_\mu \rightarrow \nu_e$  appearance.

MINOS [252] has made a combined analysis of the  $\nu_\mu$  disappearance [250] and  $\nu_\mu \rightarrow \nu_e$  appearance [41] data using the complete set of accelerator and atmospheric neutrino data. The results obtained are

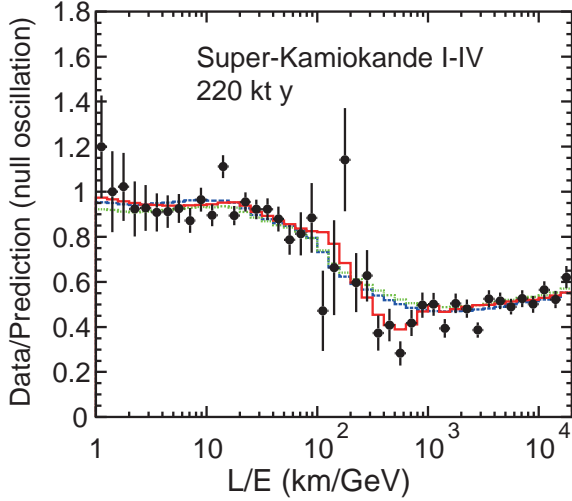
$|\Delta m_{32}^2| = (2.28 - 2.46) \times 10^{-3} \text{ eV}^2$  (68% CL) and  $\sin^2 \theta_{23} = 0.35 - 0.65$  (90% CL) for normal mass ordering and  $|\Delta m_{32}^2| = (2.32 - 2.53) \times 10^{-3} \text{ eV}^2$  (68% CL) and  $\sin^2 \theta_{23} = 0.34 - 0.67$  (90% CL) for inverted mass ordering. From this analysis, the best-fit value of  $\sin^2 \theta_{23} < 0.5$  ( $\theta_{23} < \pi/4$ ) is obtained for inverted ordering.

T2K [249] has estimated these parameters in two methods. In an analysis of  $\nu_\mu$  disappearance alone [249], the 1D 68% CL intervals obtained are  $\sin^2 \theta_{23} = 0.514_{-0.056}^{+0.055}$  and  $\Delta m_{32}^2 = (2.51 \pm 0.10) \times 10^{-3} \text{ eV}^2$  for normal mass ordering and  $\sin^2 \theta_{23} = 0.511 \pm 0.055$  and  $\Delta m_{32}^2 = (2.48 \pm 0.10) \times 10^{-3} \text{ eV}^2$  for inverted mass ordering. These results are derived by fitting the reconstructed neutrino energy spectrum of 120 1-ring  $\mu$ -like events. In an analysis of combined measurements of  $\nu_\mu$  disappearance and  $\nu_\mu \rightarrow \nu_e$  appearance with inclusion of reactor data, normal mass ordering is weakly favored, and 1D 68% CL intervals obtained are  $\sin^2 \theta_{23} = 0.528_{-0.038}^{+0.055}$  and  $\Delta m_{32}^2 = (2.51 \pm 0.11) \times 10^{-3} \text{ eV}^2$ . The T2K results for  $\sin^2 \theta_{23}$  are consistent with maximal mixing,  $\theta_{23} = \pi/4$ .

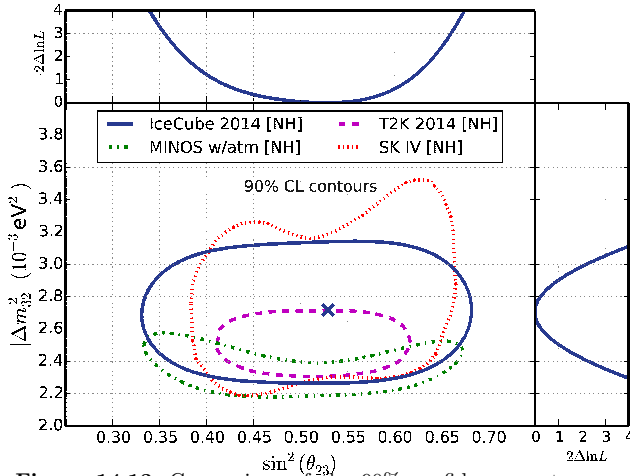
#### 14.11.3. $\bar{\nu}_\mu$ disappearance :

The CPT symmetry requires neutrinos and antineutrinos to have the same masses and mixing parameters. In vacuum, this means the same survival probabilities for a neutrino and an antineutrino which have the same energy and which traveled the same distance. In matter,  $\nu_\mu$  and  $\bar{\nu}_\mu$  survival probabilities are different, but with the experimental conditions of MINOS and T2K, the differences are small.

MINOS first observed muon antineutrino disappearance [253] with the NUMI beam line optimized for  $\bar{\nu}_\mu$  production. Actually, MINOS produced a “ $\nu_\mu$ -dominated” or “ $\bar{\nu}_\mu$ -enhanced” beam by selectively focusing positive or negative pions and kaons. In Ref. 250, MINOS reported the results of the neutrino oscillation analysis based on the data obtained with  $10.71 \times 10^{20}$  POT of the  $\nu_\mu$ -dominated beam and  $3.36 \times 10^{20}$  POT of the  $\bar{\nu}_\mu$ -enhanced beam. In addition, they used the atmospheric neutrino data based on the MINOS far detector exposure



**Figure 14.12:** Results of the L/E analysis of SK1-SK4 atmospheric neutrino data. The points show the ratio of the data to the Monte Carlo prediction without oscillations, as a function of the reconstructed L/E. The error bars are statistical only. The solid line shows the best fit with 2-flavour  $\nu_\mu \leftrightarrow \nu_\tau$  oscillations. The dashed and dotted lines show the best fit expectations for neutrino decay and neutrino decoherence hypotheses, respectively. (This figure is provided by the Super-Kamiokande Collab.)



**Figure 14.13:** Comparison of the 90% confidence contours on the atmospheric oscillation parameters derived from the T2K [249], MINOS [250], Super-Kamiokande [251], and Icecube [214] experiments. The Icecube's log-likelihood profiles for individual oscillation parameters are also shown (right and top). A normal mass ordering is assumed. This figure is taken from arXiv:1410.7227v2.

of 37.88 kt-yr [210]. Because the MINOS detector has a capability to separate neutrinos and antineutrinos on an event-by-event basis, it can use both  $\nu_\mu$  and  $\bar{\nu}_\mu$  contained events from the  $\nu_\mu$ -dominated beam. From the  $\bar{\nu}_\mu$ -enhanced beam,  $\bar{\nu}_\mu$  contained events are used. For the complete data sets used, refer to Ref. 250. Assuming the identical oscillation parameters for neutrinos and antineutrinos, the results of the fit within the two-neutrino oscillation framework using the full MINOS data sample yielded  $\Delta m^2 = (2.41^{+0.09}_{-0.10}) \times 10^{-3} \text{ eV}^2$  and  $\sin^2 2\theta = 0.950^{+0.035}_{-0.036}$ , or  $\sin^2 2\theta > 0.890$  at 90% CL. Allowing independent oscillations for neutrinos and antineutrinos, characterised respectively by  $\Delta m^2$ ,  $\theta$  and  $\bar{\Delta m}^2$ ,  $\bar{\theta}$ , the results of the fit are  $\bar{\Delta m}^2 = (2.50^{+0.23}_{-0.25}) \times 10^{-3} \text{ eV}^2$  and  $\sin^2 2\bar{\theta} = 0.97^{+0.03}_{-0.08}$ , or  $\sin^2 2\bar{\theta} > 0.83$  at 90% CL, and  $\Delta m^2 - \bar{\Delta m}^2 = (0.12^{+0.24}_{-0.26}) \times 10^{-3} \text{ eV}^2$ . This result shows that the neutrino and antineutrino mass splittings

are in agreement, which is compatible with CPT invariance.

T2K also observed  $\bar{\nu}_\mu$  disappearance [59] using an off-axis quasi-monochromatic  $\bar{\nu}_\mu$  beam peaked at  $\sim 0.6 \text{ GeV}$  with the polarity of the horn current set to focus negative pions. With  $4.01 \times 10^{20}$  POT, 34 fully contained  $\mu$ -like events were observed at the Super-Kamiokande. Oscillation parameters  $\sin^2 \theta_{23}$  and  $\Delta \bar{m}^2_{32}$  are estimated in the three-neutrino oscillation framework assuming the normal mass ordering, and all other parameters fixed to the values taken from the previous T2K fits [57] and Review of Particle Properties [103]. The best-fit parameters obtained are  $\Delta \bar{m}^2_{32} = 2.51 \times 10^{-3} \text{ eV}^2$  and  $\sin^2 \bar{\theta}_{23} = 0.45$ , with 68% confidence intervals of  $(2.26 - 2.80) \times 10^{-3} \text{ eV}^2$  for  $\Delta \bar{m}^2_{32}$  and  $0.38 - 0.64$  for  $\sin^2 \bar{\theta}_{23}$ . These results are consistent with the MINOS  $\bar{\nu}_\mu$  disappearance results [253] as well as the  $\nu_\mu$  disappearance parameters measured by T2K [57].

#### 14.11.4. $\nu_\tau$ appearance :

Super-Kamiokande Collaboration searched for the appearance of  $\tau$  leptons from the CC interactions of oscillation-generated  $\nu_\tau$  in the detector using the atmospheric neutrino data [254,24]. An excess of  $\tau$ -like events is expected in the upward-going direction. Though the Super-Kamiokande detector cannot identify a CC  $\nu_\tau$  interaction on an event by event basis, the Super-Kamiokande Collaboration excluded the no-tau-appearance hypothesis at the  $3.8\sigma$  level through a neural network analysis on the zenith-angle distribution of multi-GeV contained events [24].

For the purpose of demonstrating the appearance of tau neutrinos on an event-by-event basis, a promising method is an accelerator long-baseline experiment using emulsion technique to identify short-lived  $\tau$  leptons produced in the  $\nu_\tau$  CC interactions. OPERA adopted this strategy and searched for the appearance of  $\nu_\tau$  in the CNGS muon neutrino beam during 2008 and 2012, corresponding to a live exposure of  $17.97 \times 10^{19}$  POT in total. In 2010, OPERA reported observation of the first  $\nu_\tau$  candidate [141]. As of July 2015, OPERA has reported observation of the fifth  $\nu_\tau$  candidate [255]. The observed candidate events are classified into the four decay channels,  $\tau \rightarrow 1h$  (hadronic 1-prong),  $\tau \rightarrow 1h$  (hadronic 3-prong),  $\tau \rightarrow \mu$ , and  $\tau \rightarrow e$ , and expected signal and background events are calculated for each decay channel. The expected total signal and background events are, respectively,  $2.64 \pm 0.53$  and  $0.25 \pm 0.05$ . With 5 events observed, the OPERA Collaboration concludes the discovery of  $\nu_\tau$  appearance with a significance larger than  $5\sigma$ .

## 14.12. Measurements of $\theta_{13}$

In 2012, the last neutrino mixing angle  $\theta_{13}$  was established to be non-zero by reactor experiments. The measured value of  $\theta_{13}$  was sufficiently large to widen the opportunities to measure the unknown CP-violating phase  $\delta$  in the PMNS matrix and the neutrino mass ordering.

### 14.12.1. Overview :

Reactor  $\bar{\nu}_e$  disappearance experiments with  $L \sim 1 \text{ km}$ ,  $\langle E \rangle \sim 3 \text{ MeV}$  are sensitive to  $\sim E/L \sim 3 \times 10^{-3} \text{ eV}^2 \sim |\Delta m^2_{31(32)}|$ . At this baseline distance, the reactor  $\bar{\nu}_e$  oscillations driven by  $\Delta m^2_{21}$  are negligible. Therefore, as can be seen from Eq. (14.44) and Eq. (14.46),  $\theta_{13}$  can be directly measured. The first reactor neutrino oscillation experiment of this kind, Chooz [51] found no evidence for  $\bar{\nu}_e$  disappearance and set a 90% CL excluded region on the  $\sin^2 2\theta - \Delta m^2$  plane. Another reactor experiment Palo Verde reported somewhat less restrictive results [231].

In the accelerator neutrino oscillation experiments with conventional neutrino beams,  $\theta_{13}$  can be measured using  $\nu_\mu \rightarrow \nu_e$  appearance. In 2011, experimental indications of  $\nu_\mu \rightarrow \nu_e$  oscillations and a non-zero  $\theta_{13}$  was reported by the T2K experiment. T2K observed, with  $1.43 \times 10^{20}$  POT, six  $\nu_e$  candidate events, while the expectation for  $\theta_{13} = 0$  was  $1.5 \pm 0.3$  events. This result implied a non-zero  $\theta_{13}$  with statistical significance of  $2.5\sigma$  [28]. The MINOS Collaboration also searched for the  $\nu_\mu \rightarrow \nu_e$  appearance signal, and obtained the results that disfavored the  $\theta_{13} = 0$  hypothesis at the 89% CL [29].

In 2012, the three reactor neutrino experiments Double Chooz [30], Daya Bay [31], and RENO [32], published their first results on  $\bar{\nu}_e$

disappearance. With measurements by a far detector, Double Chooz ruled out the no-oscillation hypothesis at the 94.6% C.L. [30]. Then, Daya Bay reported  $5.2\sigma$  evidence for non-zero  $\theta_{13}$  from live-time exposure in 55 days [31]. Also, RENO reported non-zero  $\theta_{13}$  with a significance of  $4.9\sigma$  from 229 days of exposure [32]. Both Daya Bay and RENO results were obtained from rate-only analyses of the  $\bar{\nu}_e$  disappearance measurements with near and far detectors. These three experiments have been accumulating statistics and improved results have been frequently reported. For their latest results, see Section 14.12.2.

In 2014, T2K announced [42] the observation of 28  $\nu_e$  appearance events with  $4.92 \pm 0.55$  predicted background events. For  $\sin^2 2\theta_{23} = 1$  and  $\delta = 0$ , this result means that  $\theta_{13} = 0$  is excluded with a significance of  $7.3\sigma$ . For more details of the T2K results [42] as well as the 2013 MINOS results [41], see Section 14.12.3.

#### 14.12.2. Latest reactor results :

The latest Daya Bay results [38] are obtained from a total exposure of  $6.9 \times 10^5 \text{ GW}_{\text{th}} \cdot \text{ton} \cdot \text{days}$  over 404 days from October 2012 to November 2013. The Daya Bay collaboration has adopted the three-flavour oscillation scheme and analyzed the relative antineutrino rates and energy spectra between detectors using a method to predict the signal in the far hall based on measurements obtained in the near halls. With this method, they have minimized the model dependence on reactor antineutrino emission. Also, improvements in energy calibration (0.2% between detectors) and background estimation helped reduce systematic errors. Their reported new result,  $\sin^2 2\theta_{13} = 0.084 \pm 0.005$  is the most precise measurement of  $\theta_{13}$  to date. To obtain this result, they used  $\sin^2 2\theta_{12} = 0.857 \pm 0.024$  and  $\Delta m_{21}^2 = (7.50 \pm 0.20) \times 10^{-5} \text{ eV}^2$ , but the dependence on these parameters is weak. They also found for the effective mass-squared difference  $|\Delta m_{ee}^2| = (2.42 \pm 0.11) \times 10^{-3} \text{ eV}^2$ , where  $\Delta m_{ee}^2$  is obtained by replacing  $\cos^2 \theta_{12} \sin^2 \Delta m_{31}^2 L / (4E) + \sin^2 \theta_{12} \sin^2 \Delta m_{32}^2 L / (4E)$  with  $\sin^2 \Delta m_{ee}^2 L / (4E)$ . From the measured value of  $|\Delta m_{ee}^2|$ , they deduce  $\Delta m_{32}^2 = (2.37 \pm 0.11) \times 10^{-3} \text{ eV}^2$  for the normal mass ordering and  $\Delta m_{32}^2 = -(2.47 \pm 0.11) \times 10^{-3} \text{ eV}^2$  for the inverted mass ordering. These results on  $\Delta m_{32}^2$  are consistent with the T2K and MINOS results.

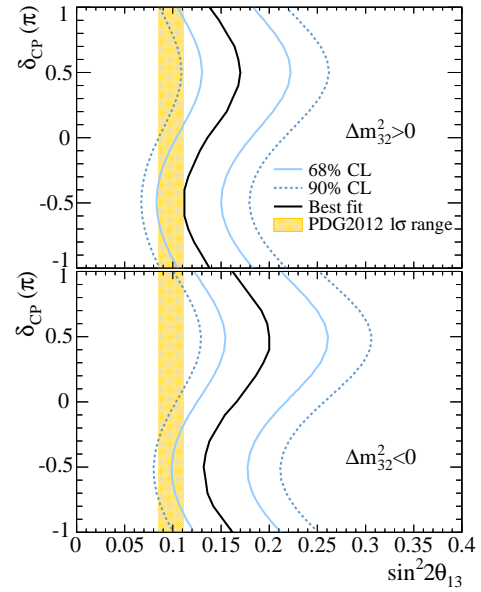
The latest RENO results are reported in 2016 [39] based on 500 live days of data. From the measured far-to-near ratio of prompt spectra,  $\sin^2 2\theta_{13} = 0.082 \pm 0.009 \pm 0.006$  and  $|\Delta m_{ee}^2| = (2.62^{+0.21+0.12}_{-0.23-0.13}) \times 10^{-3} \text{ eV}^2$  have been obtained.

The latest results from Double Chooz using the data collected by the far detector in 467.90 live days have been published in Ref. 40. From a fit to the observed spectrum (“Rate + Shape analysis”) in the two-flavour oscillation scheme, Double Chooz obtained  $\sin^2 2\theta_{13} = 0.090^{+0.032}_{-0.029}$  for the normal mass ordering using  $\Delta m_{31}^2 = (2.44^{+0.09}_{-0.10}) \times 10^{-3} \text{ eV}^2$  from MINOS [252]. For the inverted mass ordering, they obtained  $\sin^2 2\theta_{13} = 0.092^{+0.033}_{-0.029}$ , using  $|\Delta m_{31}^2| = (2.38^{+0.09}_{-0.10}) \times 10^{-3} \text{ eV}^2$  [252].

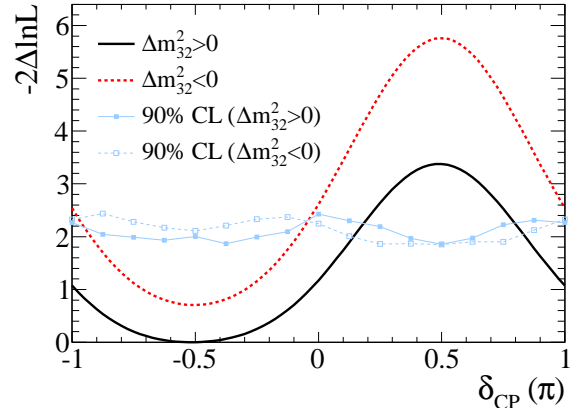
#### 14.12.3. $\nu_\mu \rightarrow \nu_e$ appearance and constraints on $\delta$ :

By examining the expression for the probability of  $\nu_\mu \rightarrow \nu_e$  oscillations in matter (given by Eq. (14.72)) it is understood that subleading terms could have rather large effects and the unknown CP-violating phase  $\delta$ , in particular, causes uncertainties in determining the value of  $\theta_{13}$ . Actually, from the measurement of  $\nu_\mu \rightarrow \nu_e$  appearance,  $\theta_{13}$  is given as a function of  $\delta$  for a given sign and value of  $\Delta m_{31}^2$ , and values of  $\theta_{23}$ ,  $\Delta m_{21}^2$  and  $\theta_{12}$ . Therefore, a single experiment with a neutrino beam cannot determine the value of  $\theta_{13}$ , although it is possible to establish a non-zero  $\theta_{13}$ . On the other hand, a combination of the  $\nu_\mu \rightarrow \nu_e$  appearance results and the precise measurements of  $\theta_{13}$  from the reactor experiments can yield constraints on the possible value of  $\delta$ .

In Ref. 42, with the observation of 28  $\nu_e$  appearance events with  $4.92 \pm 0.55$  predicted background events, for  $\delta = 0$ ,  $\sin^2 2\theta_{23} = 0.5$  and  $|\Delta m_{31(32)}^2| = 2.4 \times 10^{-3} \text{ eV}^2$ , the T2K collaborations finds in the case of  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ):  $\sin^2 2\theta_{13} = 0.140^{+0.038}_{-0.032}$  ( $0.170^{+0.045}_{-0.037}$ ). The significance of nonzero  $\theta_{13}$  is  $7.3\sigma$  for these fixed values of  $\theta_{23}$  and  $\delta$ . Changing the values of  $\theta_{23}$  and  $\delta$  within uncertainties, the



**Figure 14.14:** The T2K 68% and 90% CL allowed regions for  $\sin^2 2\theta_{13}$ , as a function of the CP violating phase  $\delta$  assuming normal mass ordering (top) and inverted mass ordering (bottom). The solid line represents the best fit  $\sin^2 2\theta_{13}$  value for given  $\delta$  values. The values of  $\sin^2 \theta_{23}$  and  $\Delta m_{32}^2$  are varied in the fit with the constraint from Ref. 23. The shaded region shows the average  $\theta_{13}$  value from Ref. 256. This figure is taken from arXiv:1311.4750.



**Figure 14.15:** The T2K value of  $-2\Delta \ln \mathcal{L}$  as a function of the CP violating phase  $\delta$  for normal mass ordering (solid line) and inverted mass ordering (dotted line). The likelihood is marginalized over  $\sin^2 2\theta_{13}$ ,  $\sin^2 \theta_{23}$  and  $\Delta m_{32}^2$ . The solid (dotted) line with markers corresponds to the 90% CL limits for normal (inverted) mass ordering, evaluated by using the Feldman-Cousins method. The  $\delta$  regions with values above the lines are excluded at 90% CL. This figure is taken from arXiv:1311.4750.

significance remains above  $7\sigma$ . The best fit value of  $\sin^2 2\theta_{13}$  thus found in the T2K experiment is approximately by a factor of 1.6 (1.9) bigger than that found in the Daya Bay experiment [36]. This implies that the compatibility of the results of the two experiments on  $\sin^2 2\theta_{13}$  requires, in particular, that  $\delta \neq 0$  and/or  $\sin^2 \theta_{23} \neq 0.5$ . As we have seen in Section 14.2, the indicated results lead to a certain indication about the possible value of  $\delta$  in the global analyses of the neutrino oscillation data. T2K has also calculated the 68% and 90% CL allowed regions for  $\sin^2 2\theta_{13}$ , as a function of  $\delta$ . The results are shown in Fig. 14.14 where the  $1\sigma$  range of the  $\sin^2 2\theta_{13}$  value from reactor experiments is also shown. T2K has further calculated



constraints on  $\delta$  by combining the  $\nu_\mu \rightarrow \nu_e$  appearance results with the  $\theta_{13}$  value from reactor experiments into a likelihood function  $\mathcal{L}$ . Fig. 14.15 shows the  $-2\Delta\ln\mathcal{L}$  value as a function of  $\delta$ . From this figure, it is seen that the combined T2K and reactor measurements prefer  $\delta = -\pi/2$  or  $3\pi/2$  for both the normal and inverted mass ordering, and some  $\delta$  regions are excluded at 90% CL.

In Ref. 41, MINOS has extended the analysis using  $10.6 \times 10^{20}$  POT  $\nu$ -beam mode and  $3.3 \times 10^{20}$  POT  $\bar{\nu}$ -beam mode appearance data. Assuming  $\Delta m_{31(32)}^2 > 0$  ( $\Delta m_{31(32)}^2 < 0$ ),  $\delta = 0$ , and  $\theta_{23} < \pi/4$ , the results of this analysis imply that  $0.01$  ( $0.03$ )  $< 2 \sin^2\theta_{23}$   $\sin^2 2\theta_{13} < 0.12$  ( $0.18$ ) at the 90% CL, with the best fit value  $2 \sin^2\theta_{23} \sin^2 2\theta_{13} = 0.051_{-0.030}^{+0.038}$  ( $0.093_{-0.049}^{+0.054}$ ). MINOS has also placed constraints on the value of  $\delta$  by combining the full MINOS appearance data and the reactor measurements of  $\theta_{13}$ .

In 2016, NO $\nu$ A reported its first result of  $\nu_e$  appearance measurement [56]. The data were taken between February 2014 and May 2015. During this period, the effective fiducial mass of the NO $\nu$ A far detector varied from 2.3 kt for 4.0 kt of total mass to 10 kt for the full 14 kt. The exposure corresponding to these data is equivalent to  $2.74 \times 10^{20}$  POT collected in the full 14 kt far detector, and 6  $\nu_e$ -like events were observed compared to  $0.99 \pm 0.11$  (syst.) expected background events: a  $3.3\sigma$  excess over the background prediction. NO $\nu$ A also quote the result of a secondary event selection method. With this analysis, 11 events were observed over the expected background of  $1.07 \pm 0.14$  (syst.) events.

### 14.13. Search for Oscillations Involving Light Sterile Neutrinos

Although the mixing of the 3 flavour neutrino states has been experimentally well established, implying the existence of 3 light neutrinos  $\nu_j$  having masses  $m_j$  not exceeding approximately 1 eV, there have been possible hints for the presence in the mixing of one or more additional neutrino states with masses at the eV scale. If these states exist, they must be related to the existence of one or more sterile neutrinos (sterile neutrino fields) which mix with the active flavour neutrinos (active flavour neutrino fields). The hints under discussion have been obtained: i) in the LSND  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance experiment [225], in which a significant excess of events over the background is claimed to have been observed, ii) from the analysis of the  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  [228] and  $\nu_\mu \rightarrow \nu_e$  [227] appearance data of the MiniBooNE experiment, iii) from the re-analyses of the short baseline (SBL) reactor neutrino oscillation data using newly calculated fluxes of reactor  $\bar{\nu}_e$  [257,119], which show a possible “disappearance” of the reactor  $\bar{\nu}_e$  (“reactor neutrino anomaly”), and iv) from the data of the radioactive source calibrations of the GALLEX [202] and SAGE [203] solar neutrino experiments.

The short baseline neutrino oscillation experiment MiniBooNE at Fermilab investigated  $\nu_e$  [227] and  $\bar{\nu}_e$  [228] appearance in  $\nu_\mu$  and  $\bar{\nu}_\mu$  beams, respectively, with a detector containing 800 tons of mineral oil and located 541 m downstream of the production target. With the antineutrino running mode [228], a  $2.8\sigma$  excess of events over the background was observed in the energy range of  $200 < E_\nu < 1250$  MeV in the charged-current quasielastic data. Excess events were observed, in particular, in the interval of energies  $200 < E_\nu < 475$  MeV, which corresponds to  $L/E$  range outside of that probed in the LSND experiment. The origin of this excess is not understood. Employing a simple 2-neutrino oscillation hypothesis and using the data from the entire neutrino energy interval  $200 < E_\nu < 1250$  MeV used in the data analysis, this result, interpreted in terms of  $\nu_\mu \rightarrow \nu_e$  oscillations, corresponds to an allowed region in the  $\sin^2 2\theta - \Delta m^2$  plane, which overlaps with the allowed region obtained from the interpretation of the LSND data in terms of  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations. The overlap region at the 90% CL extends over  $\Delta m^2 \sim \text{a few } \times 10^{-2} \text{ eV}^2$  at  $\sin^2 2\theta = 1$  to  $1 \text{ eV}^2$  at  $\sin^2 2\theta = \text{a few } \times 10^{-3}$ . The MiniBooNE Collaboration studied also the CP conjugate oscillation channel [227],  $\nu_\mu \rightarrow \nu_e$ , and observed a  $3.4 \sigma$  excess of events in the same energy range. Most of the excess events lie in the interval  $200 < E_\nu < 475$  MeV and are incompatible with the  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation interpretation of the LSND data. The energy spectra of the excess events observed

in the  $\nu_\mu$  [227] and  $\bar{\nu}_\mu$  [228] runs are only marginally compatible with each other and thus with the simple 2-neutrino oscillation hypothesis.

The reactor neutrino anomaly [257] is related to the results of a new and very detailed calculation of the reactor  $\bar{\nu}_e$  fluxes [119] which were found to be by approximately 3.5% larger than the fluxes calculated in Ref. 122 and widely used in the past in the interpretation of the data of the SBL reactor  $\bar{\nu}_e$  oscillation experiments. These data show indications for reactor  $\bar{\nu}_e$  “disappearance” when analyzed using the fluxes from [119]. It should be added that there are a number of uncertainties in the calculation of the fluxes under discussion (associated, *e.g.*, with the weak magnetism term contribution to the corresponding  $\beta$ -decay rates [120], the contribution of a relatively large number of “forbidden”  $\beta$ -decays [258], etc.) which can be of the order of the difference between the “old” and “new” fluxes.

Radioactive source calibrations of the GALLEX [202] and SAGE [203] experiments also showed a deficit of the measured fluxes compared to the expected fluxes (“Gallium anomaly”), and therefore might be interpreted as hints for  $\nu_e$  disappearance.

Significant constraints on the parameters characterizing the oscillations involving sterile neutrinos follow from the negative results of the searches for  $\nu_\mu \rightarrow \nu_e$  and/or  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations in the Karmen [226], NOMAD [259], ICARUS [224], and OPERA [260] experiments, and from the nonobservation of effects of oscillations into sterile neutrinos in the solar neutrino experiments and in the studies of  $\nu_\mu$  and/or  $\bar{\nu}_\mu$  disappearance in the CDHSW [261], MINOS and SuperKamiokande experiments.

Two possible “minimal” phenomenological models (or schemes) with light sterile neutrinos are widely used in order to explain the data discussed in this section in terms of neutrino oscillations: the so-called “3 + 1” and “3 + 2” models. They contain respectively one and two sterile neutrinos (right-handed sterile neutrino fields). Thus, the “3 + 1” and “3 + 2” models have altogether 4 and 5 light massive neutrinos  $\nu_j$ , which in the minimal versions of these models are Majorana particles. The additional neutrinos  $\nu_4$  and  $\nu_5$  should have masses  $m_4$  and  $m_5$  at the eV scale (see below). It follows from the data that if  $\nu_4$  or  $\nu_5$  exist, they couple to the electron and muon in the weak charged lepton current with couplings  $U_{ek}$  and  $U_{\mu k}$ ,  $k = 4, 5$ , which are approximately  $|U_{ek}| \sim 0.1$  and  $|U_{\mu k}| \sim 0.1$ .

Global analysis of all the data (positive evidences and negative results) relevant for the test of the sterile neutrino hypothesis were performed in Ref. 262 and in Ref. 263. Analyzing the data within the 3 + 1 scheme, the authors of Ref. 262 find for the best fit values of the parameters  $|U_{e4}|^2$ ,  $|U_{\mu 4}|^2$  and  $\Delta m_{\text{SBL}}^2 \equiv m_4^2 - m_{\text{min}}^2$ , where  $m_{\text{min}} = \min(m_j)$ ,  $j = 1, 2, 3$ , characterizing the active-sterile neutrino (antineutrino) oscillations:

$$|U_{e4}|^2 = 0.0225, \quad |U_{\mu 4}|^2 = 0.0289, \quad \Delta m_{\text{SBL}}^2 = 0.93 \text{ eV}^2. \quad (14.102)$$

In contrast to Ref. 262, the authors of Ref. 263 reported also results within the 3 + 1 scheme without including in the data set used in their global analysis the MiniBooNE data at  $E_\nu \leq 0.475$  GeV. As we have already mentioned, these data show an excess of events over the estimated background [227,228] whose nature is presently not well understood. For the best fit values of  $|U_{e4}|^2$ ,  $|U_{\mu 4}|^2$  and  $\Delta m_{\text{SBL}}^2$  in this case the authors of Ref. 263 find:

$$|U_{e4}|^2 = 0.03, \quad |U_{\mu 4}|^2 = 0.013, \quad \Delta m_{\text{SBL}}^2 = 1.60 \text{ eV}^2. \quad (14.103)$$

In the context of the “3+1” model, the Daya Bay Collaboration recently searched for relative spectral distortion in their reactor antineutrino data, due to possible mixing of a light sterile neutrino in the  $|\Delta m_{41}^2| < 0.3 \text{ eV}^2$  region [264]. The result is consistent with no sterile neutrino mixing, leading to the most stringent limits on  $\sin^2\theta_{14}$  in the  $10^{-3} \text{ eV}^2 < |\Delta m_{41}^2| < 0.1 \text{ eV}^2$  region.

The existence of light sterile neutrinos has cosmological implications the discussion of which lies outside the scope of the present article (for a discussion of the cosmological constraints on light sterile neutrinos see, *e.g.*, [265,75]).

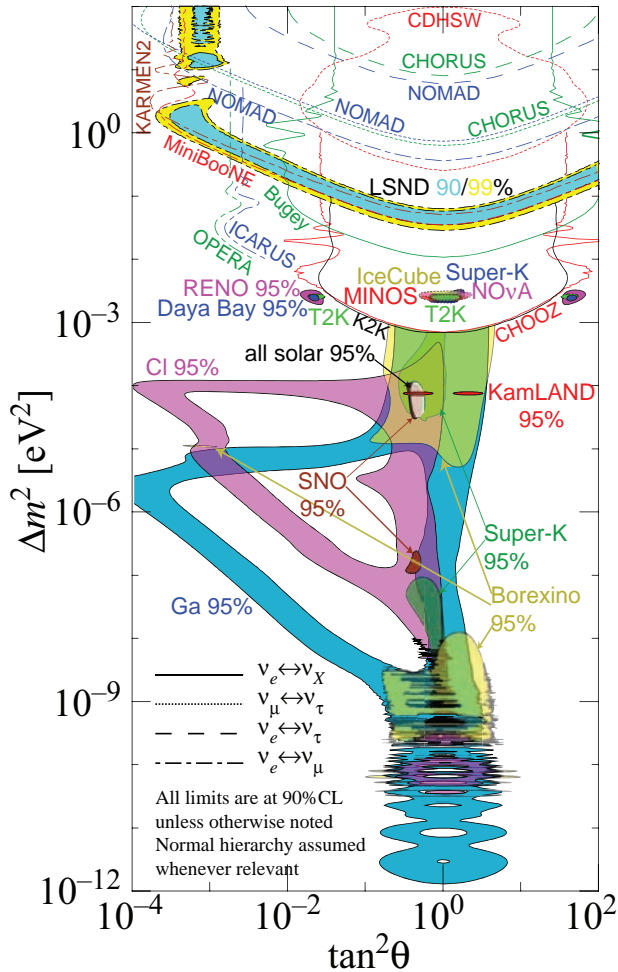
The hypothesis of existence of light sterile neutrinos with eV scale masses and charged current couplings to the electron and muon quoted

above will be tested in a number of experiments with reactor and accelerator neutrinos, and neutrinos from artificial sources, some of which are under preparation and planned to start taking data already this year (see, *e.g.*, [266] for a detailed list and discussion of the planned experiments).

#### 14.14. Outlook

The currently available data on neutrino oscillations are summarised in Fig. 14.16.

The program of experimental research in neutrino physics extends beyond 2030 (see, *e.g.*, Refs. [80,83,84,85,219]). In the coming years we expect a wealth of new data that, it is hoped, will shed light on the fundamental aspects of neutrino mixing: the nature - Dirac or Majorana - of massive neutrinos, the type of spectrum the neutrino masses obey, the status of CP symmetry in the lepton sector, the absolute neutrino mass scale, the origin of the observed patterns of the neutrino masses and mixing, and, eventually, on the mechanism of neutrino mass generation. We are looking forward to these exciting developments in neutrino physics.



**Figure 14.16:** The regions of squared-mass splitting and mixing angle favored or excluded by various neutrino oscillation experiments. The figure was contributed by H. Murayama (University of California, Berkeley, and Kavli IPMU, University of Tokyo). References to the data used in the figure and the description of how the figure was obtained can be found at <http://hitoshi.berkeley.edu/neutrino>.

#### References:

1. B. Pontecorvo, Zh. Eksp. Teor. Fiz. **53**, 1717 (1967) [Sov. Phys. JETP **26**, 984 (1968)].
2. M. Fukugita and T. Yanagida, Phys. Lett. **B174**, 45 (1986); V.A. Kuzmin, V.A. Rubakov, and M.E. Shaposhnikov, Phys. Lett. **B155**, 36 (1985).
3. P. Minkowski, Phys. Lett. **B67**, 421 (1977); see also: M. Gell-Mann, P. Ramond, and R. Slansky in *Supergravity*, p. 315, edited by F. Nieuwenhuizen and D. Friedman, North Holland, Amsterdam, 1979; T. Yanagida, *Proc. of the Workshop on Unified Theories and the Baryon Number of the Universe*, edited by O. Sawada and A. Sugamoto, KEK, Japan 1979; R.N. Mohapatra and G. Senjanović, Phys. Rev. Lett. **44**, 912 (1980).
4. B. Pontecorvo, Zh. Eksp. Teor. Fiz. **33**, 549 (1957) and **34**, 247 (1958).
5. Z. Maki, M. Nakagawa, and S. Sakata, Prog. Theor. Phys. **28**, 870 (1962).
6. B.T. Cleveland *et al.*, Astrophys. J. **496**, 505 (1988).
7. Y. Fukuda *et al.*, [Kamiokande Collab.], Phys. Rev. Lett. **77**, 1683 (1996).
8. J.N. Abdurashitov *et al.*, [SAGE Collab.], Phys. Rev. **C80**, 015807 (2009).
9. P. Anselmann *et al.*, [GALLEX Collab.], Phys. Lett. **B285**, 376 (1992).
10. W. Hampel *et al.*, [GALLEX Collab.], Phys. Lett. **B447**, 127 (1999).
11. M. Altmann *et al.*, [GNO Collab.], Phys. Lett. **B616**, 174 (2005).
12. S. Fukuda *et al.*, [Super-Kamiokande Collab.], Phys. Lett. **B539**, 179 (2002).
13. Q.R. Ahmad *et al.*, [SNO Collab.], Phys. Rev. Lett. **87**, 071301 (2001).
14. Q.R. Ahmad *et al.*, [SNO Collab.], Phys. Rev. Lett. **89**, 011301 (2002).
15. K. Eguchi *et al.*, [KamLAND Collab.], Phys. Rev. Lett. **90**, 021802 (2003).
16. T. Araki *et al.*, [KamLAND Collab.], Phys. Rev. Lett. **94**, 081801 (2005).
17. Y. Fukuda *et al.*, [Super-Kamiokande Collab.], Phys. Rev. Lett. **81**, 1562 (1998).
18. Y. Ashie *et al.*, [Super-Kamiokande Collab.], Phys. Rev. Lett. **93**, 101801 (2004).
19. M.H. Ahn *et al.*, [K2K Collab.], Phys. Rev. **D74**, 072003 (2006).
20. D.G. Michael *et al.*, [MINOS Collab.], Phys. Rev. Lett. **97**, 191801 (2006); P. Adamson *et al.*, [MINOS Collab.], Phys. Rev. Lett. **101**, 131802 (2008).
21. P. Adamson *et al.*, [MINOS Collab.], Phys. Rev. Lett. **106**, 181801 (2011).
22. K. Abe *et al.*, [T2K Collab.], Phys. Rev. **D85**, 031103 (2012).
23. K. Abe *et al.*, [T2K Collab.], Phys. Rev. Lett. **111**, 211803 (2013).
24. K. Abe *et al.*, [Super-Kamiokande Collab.], Phys. Rev. Lett. **110**, 181802 (2013).
25. A. Pastore, talk at the EPS HEP 2013 Conference, July 18-24, 2013, Stockholm.
26. L. Wolfenstein, Phys. Rev. **D17**, 2369 (1978); *Proc. of the 8th International Conference on Neutrino Physics and Astrophysics - "Neutrino'78"* (ed. E.C. Fowler, Purdue University Press, West Lafayette, 1978), p. C3.
27. S.P. Mikheev and A.Y. Smirnov, Sov. J. Nucl. Phys. **42**, 913 (1985); Nuovo Cimento **9C**, 17 (1986).
28. K. Abe *et al.*, [T2K Collab.], Phys. Rev. Lett. **107**, 041801 (2011).
29. P. Adamson *et al.*, [MINOS Collab.], Phys. Rev. Lett. **107**, 181802 (2011).
30. Y. Abe *et al.*, [Double Chooz Collab.], Phys. Rev. Lett. **108**, 131801 (2012).
31. F.P. An *et al.*, [Daya Bay Collab.], Phys. Rev. Lett. **108**, 171803 (2012).

32. J.K. Ahn *et al.*, [RENO Collab.], Phys. Rev. Lett. **108**, 191802 (2012).
33. Y. Abe *et al.*, [Double Chooz Collab.], Phys. Rev. **D86**, 052008 (2012).
34. K. Abe *et al.*, [T2K Collab.], Phys. Rev. **D88**, 032002 (2013).
35. F.P. An *et al.*, [Daya Bay Collab.], Chin. Phys. C **37**, 011001 (2013).
36. F.P. An *et al.*, [Daya Bay Collab.], Phys. Rev. Lett. **112**, 061801 (2014).
37. S.-H. Seo [for the RENO Collab.], talk at the TAUP2013 International Workshop, September 9-13, 2013, Asilomar, California, USA.
38. F.P. An *et al.*, [Daya Bay Collab.], Phys. Rev. Lett. **115**, 111802 (2015).
39. J.H. Choi *et al.*, [RENO Collab.], Phys. Rev. Lett. **116**, 211801 (2016).
40. Y. Abe *et al.*, [Double Chooz Collab.], JHEP **1014**, 086 (2014).
41. P. Adamson *et al.*, [MINOS Collab.], Phys. Rev. Lett. **110**, 171801 (2013).
42. K. Abe *et al.*, [T2K Collab.], Phys. Rev. Lett. **112**, 061802 (2014).
43. D. Karlen in RPP2012 [Phys. Rev. **D86**, Part I, 629 (2012)].
44. E. Majorana, Nuovo Cimento **5**, 171 (1937).
45. Majorana particles, in contrast to Dirac fermions, are their own antiparticles. An electrically charged particle (like the electron) cannot coincide with its antiparticle (the positron) which carries the opposite non-zero electric charge.
46. S.M. Bilenky and S.T. Petcov, Rev. Mod. Phys. **59**, 671 (1987).
47. S.T. Petcov, Adv. High Energy Phys. **2013**, 852987 (2013) and arXiv:1303.5819.
48. S.M. Bilenky, J. Hosek, and S.T. Petcov, Phys. Lett. **B94**, 495 (1980).
49. J. Schechter and J.W.F. Valle, Phys. Rev. **D22**, 2227 (1980); M. Doi *et al.*, Phys. Lett. **B102**, 323 (1981).
50. L. Wolfenstein, Phys. Lett. **B107**, 77 (1981); J. Bernabeu and P. Pascual, Nucl. Phys. **B228**, 21 (1983); S.M. Bilenky, N.P. Nedelcheva, and S.T. Petcov, Nucl. Phys. **B247**, 61 (1984); B. Kayser, Phys. Rev. **D30**, 1023 (1984).
51. M. Apollonio *et al.*, [Chooz Collab.], Phys. Lett. **B466**, 415 (1999); Eur. Phys. J. **C27**, 331 (2003).
52. F. Capozzi *et al.*, Phys. Rev. **D89**, 093018 (2014).
53. M.C. Gonzalez-Garcia, M. Maltoni, and T. Schwetz, JHEP **11**, 052 (2014).
54. D.V. Forero, M. Tortola and J.W.F. Valle, Phys. Rev. **D90**, 093006 (2014).
55. P. Adamson *et al.*, [NO $\nu$ A Collab.], Phys. Rev. **D93**, 051104 (2016).
56. P. Adamson *et al.*, [NO $\nu$ A Collab.] Phys. Rev. Lett. **116**, 151806 (2016).
57. K. Abe *et al.*, [T2K Collab.] Phys. Rev. **D91**, 072010 (2015).
58. M.R. Salzgeber *et al.*, [T2K Collab.], arXiv:1508.06153.
59. K. Abe *et al.*, [T2K Collab.], Phys. Rev. Lett. **116**, 181801 (2016).
60. F. Capozzi *et al.*, arXiv:1601.07777.
61. N. Cabibbo, Phys. Lett. **B72**, 333 (1978).
62. V. Barger *et al.*, Phys. Rev. Lett. **45**, 2084 (1980).
63. P.I. Krastev and S.T. Petcov, Phys. Lett. **B205**, 84 (1988).
64. C. Jarlskog, Z. Phys. **C29**, 491 (1985).
65. P. Langacker *et al.*, Nucl. Phys. **B282**, 589 (1987).
66. S.M. Bilenky, S. Pascoli, and S.T. Petcov, Phys. Rev. **D64**, 053010 (2001), and *ibid.*, 113003.
67. S. Pascoli, S.T. Petcov, and A. Riotto, Phys. Rev. **D75**, 083511 (2007); E. Molinaro and S.T. Petcov, Phys. Lett. **B671**, 60 (2009).
68. S. Pascoli, S.T. Petcov, and A. Riotto, Nucl. Phys. **B774**, 1 (2007).
69. In the convention we use, the neutrino masses are not ordered in magnitude according to their index number:  $\Delta m_{31}^2 < 0$  corresponds to  $m_3 < m_1 < m_2$ . We can also number the massive neutrinos in such a way that one always has  $m_1 < m_2 < m_3$ , see, *e.g.*, Ref. 66.
70. F. Perrin, Comptes Rendus **197**, 868 (1933); E. Fermi, Nuovo Cimento **11**, 1 (1934).
71. V. Lobashev *et al.*, Nucl. Phys. **A719**, 153c, (2003).
72. Ch. Kraus *et al.*, Eur. Phys. J. **C40**, 447 (2005).
73. K. Eitel *et al.*, Nucl. Phys. (Proc. Supp.) **B143**, 197 (2005).
74. V.N. Aseev *et al.*, Phys. Rev. **D84**, 112003 (2011).
75. O. Lahav and A.R. Liddle in RPP2014.
76. K.N. Abazajian *et al.*, Astropart. Phys. **35**, 177 (2011).
77. P.A.R. Ade *et al.*, [Planck Collab.], Astron. Astrophys. **571**, A16 (2014).
78. P.A.R. Ade *et al.*, [Planck Collab.], arXiv:1502.01589.
79. J. Bernabeu *et al.*, arXiv:1005.3146.
80. A. de Gouvea *et al.*, arXiv:1310.4340.
81. S.T. Petcov and M. Piai, Phys. Lett. **B533**, 94 (2002); S. Choubey, S.T. Petcov and M. Piai, Phys. Rev. **D68**, 113006 (2003); J. Learned *et al.*, Phys. Rev. **D78**, 071302 (2008); L. Zhan *et al.*, Phys. Rev. **D78**, 111103 (2008) and Phys. Rev. **D79**, 073007 (2009); P. Ghoshal and S.T. Petcov, JHEP **1103**, 058 (2011).
82. R.N. Cahn *et al.*, arXiv:1307.5487.
83. S.K. Agarwalla *et al.*, arXiv:1312.6520; C. Adams *et al.*, arXiv:1307.7335.
84. R. Acciarri *et al.*, [DUNE Collab.] arXiv:1601.05471 and arXiv:1601.02984.
85. K. Abe *et al.*, Prog. Theor. Exp. Phys. **5**, 053C02 (2015), arXiv:1502.05199.
86. R. Mohapatra *et al.*, Rept. on Prog. in Phys. **70**, 1757 (2007); A. Bandyopadhyay *et al.*, Rept. on Prog. in Phys. **72**, 106201 (2009).
87. G. Altarelli and F. Feruglio, Rev. Mod. Phys. **82**, 2701 (2010); S.F. King *et al.*, New J. Phys. **16**, 045018 (2014); S.T. Petcov, Phys. Lett. **B110**, 245 (1982); R. Barbieri *et al.*, JHEP **9812**, 017 (1998); P.H. Frampton, S.T. Petcov, and W. Rodejohann, Nucl. Phys. **B687**, 31 (2004).
88. A. Morales and J. Morales, Nucl. Phys. (Proc. Supp.) **B114**, 141 (2003); C. Aalseth *et al.*, hep-ph/0412300; A. Giuliani and A. Poves, Adv. High Energy Phys. **2012**, 857016 (2012) (<http://www.hindawi.com/journals/ahp/si/437630/>); J. D. Vergados, H. Ejiri and F. Simkovic, Rept. Prog. Phys. **75**, 106301 (2012).
89. S. Pascoli and S.T. Petcov, Phys. Lett. **B544**, 239 (2002); see also: S. Pascoli, S.T. Petcov, and L. Wolfenstein, Phys. Lett. **B524**, 319 (2002).
90. S.M. Bilenky *et al.*, Phys. Rev. **D54**, 4432 (1996).
91. S.M. Bilenky *et al.*, Phys. Lett. **B465**, 193 (1999); F. Vissani, JHEP **9906**, 022 (1999); K. Matsuda *et al.*, Phys. Rev. **D62**, 093001 (2000); K. Czakon *et al.*, hep-ph/0003161; H.V. Klapdor-Kleingrothaus, H. Päs and A.Yu. Smirnov, Phys. Rev. **D63**, 073005 (2001); S. Pascoli, S.T. Petcov and W. Rodejohann, Phys. Lett. **B549**, 177 (2002), and *ibid.* **B558**, 141 (2003); H. Murayama and Peña-Garay, Phys. Rev. **D69**, 031301 (2004); S. Pascoli, S.T. Petcov, and T. Schwetz, Nucl. Phys. **B734**, 24 (2006); M. Lindner, A. Merle, and W. Rodejohann, Phys. Rev. **D73**, 053005 (2006); A. Faessler *et al.*, Phys. Rev. **D79**, 053001 (2009); S. Pascoli and S.T. Petcov, Phys. Rev. **D77**, 113003 (2008); W. Rodejohann, Int. J. Mod. Phys. E **20**, 1833 (2011).
92. A. Halprin *et al.*, Phys. Rev. **D13**, 2567 (1976); H. Päs *et al.*, Phys. Lett. **B453**, 194 (1999), and Phys. Lett. **B498**, 35 (2001); F.F. Deppisch, M. Hirsch and H. Päs, J. Phys. **G39**, 124007 (2012); L.C. Helo *et al.*, JHEP **1505**, 092 (2015).
93. A. Halprin, S.T. Petcov and S.P. Rosen, Phys. Lett. **B125**, 335 (1983); F. Deppisch and H. Päs, Phys. Rev. Lett. **98**, 232501 (2007); V.M. Gehman and S.R. Elliott, J. Phys. **G34**, 667 (2007), Erratum J. Phys. **G35**, 029701 (2008).
94. A. Faessler *et al.*, Phys. Rev. **D83**, 113003 (2011); A. Faessler *et al.*, Phys. Rev. **D83**, 113015 (2011); F. Simkovic, J. Vergados, and A. Faessler, Phys. Rev. **D82**, 113015 (2010); A. Meroni, S.T. Petcov, and F. Simkovic, JHEP **1302**, 025 (2013).
95. For alternative mechanisms of neutrino mass generation see, *e.g.*, the first article in Ref. 86 and references quoted therein.

96. S. Davidson and A. Ibarra, Phys. Lett. **B535**, 25 (2002).
97. A. Abada *et al.*, JCAP **0604**, 004 (2006); E. Nardi *et al.*, JHEP **0601**, 164 (2006).
98. J.N. Bahcall, A.M. Serenelli, and S. Basu, Astrophys. J. **621**, L85 (2005).
99. J.N. Bahcall, A.M. Serenelli, and S. Basu, Astrophys. J. Supp. **165**, 400 (2006).
100. C. Peña-Garay and A.M. Serenelli, arXiv:0811.2424.
101. L.C. Stonehill, J.A. Formaggio, and R.G.H. Robertson, Phys. Rev. **C69**, 015801 (2004).
102. A.M. Serenelli, W.C. Haxton, and C. Peña-Garay, Astrophys. J. **743**, 24 (2011).
103. K.A. Olive *et al.*, (PDG), Chin. Phys. C **38**, 090001 (2014).
104. M. Honda *et al.*, Phys. Rev. **D70**, 043008 (2004); Phys. Rev. **D75**, 043006 (2007); Phys. Rev. **D83**, 123001 (2011).
105. M. Honda *et al.*, Phys. Rev. **D92**, 023004 (2015).
106. G.D. Barr *et al.*, Phys. Rev. **D70**, 023006 (2004).
107. G. Battistoni *et al.*, Astropart. Phys. **19**, 269 (2003).
108. S.E. Kopp, Phys. Rep. **439**, 101 (2007).
109. A. Ambrosini *et al.*, Phys. Lett. **B420**, 225 (1998); Phys. Lett. **B425**, 208 (1998).
110. M.G. Catanesi *et al.*, Nucl. Instrum. Methods **A571**, 527 (2007).
111. R. Raja, Nucl. Instrum. Methods **A553**, 225 (2005).
112. N. Abgrall *et al.*, J. Instr. **9**, P06005 (2014).
113. D. Beavis *et al.*, Physics Design Report, BNL 52459 (1995).
114. K. Abe *et al.*, [T2K Collab.], Phys. Rev. **D87**, 012001 (2013).
115. R.B. Patterson [NOvA Collab.], Nucl. Phys. (Proc. Supp.) **B235-236**, 151 (2013).
116. K. Schrenkenbach, ILL technical report 84SC26T, quoted in M. Apollonio *et al.*, Eur. Phys. J. **C27**, 331 (2003).
117. J. Cao, Nucl. Phys. (Proc. Supp.) **B229-232**, 205 (2012).
118. M.F. James, J. Nucl. Energy **23**, 517 (1969); V. Kopeikin *et al.*, Phys. Atom. Nucl. **67**, 1892 (2004).
119. T.A. Mueller *et al.*, Phys. Rev. **C83**, 054615 (2011).
120. P. Huber, Phys. Rev. **C84**, 024617 (2011).
121. F. von Feilitzsch *et al.*, Phys. Lett. **B118**, 162 (1982).
122. K. Schreckenbach *et al.*, Phys. Lett. **B160**, 325 (1985).
123. A.A. Hahn *et al.*, Phys. Lett. **B218**, 365 (1989).
124. P. Vogel *et al.*, Phys. Rev. **C24**, 1543 (1981).
125. N. Haag *et al.*, Phys. Rev. Lett. **112**, 122501 (2014).
126. F.P. An *et al.*, [Daya Bay Collab.], Phys. Rev. Lett. **116**, 061801 (2016).
127. S.-B. Kim, arXiv:1412.2199.
128. D.A. Dwyer and T.J. Langford, Phys. Rev. Lett. **114**, 012502 (2014); A.A. Sonzogni, T.D. Johnson, and E.A. McCutchan, Phys. Rev. **C91**, 011301R (2015); A.C. Hayes *et al.*, Phys. Rev. **D92**, 033015 (2015).
129. C. Bemporad, G. Gratta, and P. Vogel, Rev. Mod. Phys. **74**, 292 (2002).
130. S. Nussinov, Phys. Lett. **B63**, 201 (1976); B. Kayser, Phys. Rev. **D24**, 110 (1981); J. Rich, Phys. Rev. **D48**, 4318 (1993); H. Lipkin, Phys. Lett. **B348**, 604 (1995); W. Grimus and P. Stockinger, Phys. Rev. **D54**, 3414 (1996); L. Stodolski, Phys. Rev. **D58**, 036006 (1998); W. Grimus, P. Stockinger, and S. Mohanty, Phys. Rev. **D59**, 013011 (1999); L.B. Okun, Surv. High Energy Physics **15**, 75 (2000); J.-M. Levy, hep-ph/0004221 and arXiv:0901.0408; A.D. Dolgov, Phys. Reports **370**, 333 (2002); C. Giunti, Phys. Scripta **67**, 29 (2003) and Phys. Lett. **B17**, 103 (2004); M. Beuthe, Phys. Reports **375**, 105 (2003); H. Lipkin, Phys. Lett. **B642**, 366 (2006); S.M. Bilenky, F. von Feilitzsch, and W. Potzel, J. Phys. **G34**, 987 (2007); C. Giunti and C.W. Kim, *Fundamentals of Neutrino Physics and Astrophysics* (Oxford University Press, Oxford, 2007); E.Kh. Akhmedov, J. Kopp, and M. Lindner, JHEP **0805**, 005 (2008); E.Kh. Akhmedov and A.Yu. Smirnov, Phys. Atom. Nucl. **72**, 1363 (2009).
131. For the subtleties involved in the step leading from Eq. (14.1) to Eq. (14.26) see, *e.g.*, Ref. 132.
132. A.G. Cohen, S.L. Glashow, and Z. Ligeti, Phys. Lett. **B678**, 191 (2009).
133. The neutrino masses do not exceed approximately 1 eV,  $m_j \lesssim 1$ , while in neutrino oscillation experiments neutrinos with energy  $E \gtrsim 100$  keV are detected.
134. E. K. Akhmedov, J. Kopp and M. Lindner, JHEP **0805**, 005 (2008).
135. S.M. Bilenky and B. Pontecorvo, Phys. Reports **41**, 225 (1978).
136. In Eq. (14.31) we have neglected the possible instability of neutrinos  $\nu_j$ . In most theoretical models with nonzero neutrino masses and neutrino mixing, the predicted half life-time of neutrinos with mass of 1 eV exceeds the age of the Universe, see, *e.g.*, S.T. Petcov, Yad. Fiz. **25**, 641 (1977), (E) *ibid.*, **25** (1977) 1336 [Sov. J. Nucl. Phys. **25**, 340 (1977), (E) *ibid.*, **25**, (1977), 698], and Phys. Lett. **B115**, 401 (1982); W. Marciano and A.I. Sanda, Phys. Lett. **B67**, 303 (1977); P. Pal and L. Wolfenstein, Phys. Rev. **D25**, 766 (1982).
137. L.B. Okun (2000), J.-M. Levy (2000) and H. Lipkin (2006) quoted in Ref. 130 and Ref. 132.
138. The articles by L. Stodolsky (1998) and H. Lipkin (1995) quoted in Ref. 130.
139. V. Gribov and B. Pontecorvo, Phys. Lett. **B28**, 493 (1969).
140. A. De Rujula *et al.*, Nucl. Phys. **B168**, 54 (1980).
141. N. Agafonova *et al.*, [OPERA Collab.], Phys. Lett. **B691**, 138 (2010); New J. Phys. **14**, 033017 (2012).
142. N. Agafonova *et al.*, [OPERA Collab.], JHEP **1311**, 036 (2013).
143. S. Goswami *et al.*, Nucl. Phys. (Proc. Supp.) **B143**, 121 (2005).
144. These processes are important, however, for the supernova neutrinos see, *e.g.*, G. Raffelt, *Proc. International School of Physics "Enrico Fermi", CLII Course "Neutrino Physics"*, 23 July-2 August 2002, Varenna, Italy [hep-ph/0208024], and articles quoted therein.
145. We standardly assume that the weak interaction of the flavour neutrinos  $\nu_l$  and antineutrinos  $\bar{\nu}_l$  is described by the Standard Model (for alternatives see, *e.g.*, Ref. 26; M.M. Guzzo *et al.*, Phys. Lett. **B260**, 154 (1991); E. Roulet, Phys. Rev. **D44**, R935 (1991) and Ref. 86).
146. V. Barger *et al.*, Phys. Rev. **D22**, 2718 (1980).
147. P. Langacker, J.P. Leveille, and J. Sheiman, Phys. Rev. **D27**, 1228 (1983).
148. The difference between the  $\nu_\mu$  and  $\nu_\tau$  indices of refraction arises at one-loop level and can be relevant for the  $\nu_\mu - \nu_\tau$  oscillations in very dense media, like the core of supernovae, *etc.*; see F.J. Botella, C.S. Lim, and W.J. Marciano, Phys. Rev. **D35**, 896 (1987).
149. The relevant formulae for the oscillations between the  $\nu_e$  and a sterile neutrino  $\nu_s$ ,  $\nu_e \leftrightarrow \nu_s$ , can be obtained from those derived for the case of  $\nu_e \leftrightarrow \nu_{\mu(\tau)}$  oscillations by Refs. [65,147] replacing  $N_e$  with  $(N_e - 1/2N_n)$ ,  $N_n$  being the neutron number density in matter.
150. T.K. Kuo and J. Pantaleone, Phys. Lett. **B198**, 406 (1987).
151. A.D. Dziewonski and D.L. Anderson, Physics of the Earth and Planetary Interiors **25**, 297 (1981).
152. The first studies of the effects of Earth matter on the oscillations of neutrinos were performed numerically in Refs. [146,153] and in E.D. Carlson, Phys. Rev. **D34**, 1454 (1986); A. Dar *et al.*, *ibid.*, **D35**, 3607 (1988); in Ref. 63 and in G. Auriemma *et al.*, *ibid.*, **D37**, 665 (1988).
153. A.Yu. Smirnov and S.P. Mikheev, *Proc. of the VIth Moriond Workshop* (eds. O. Fackler, J. Tran Thanh Van, Frontières, Gif-sur-Yvette, 1986), p. 355.
154. S.T. Petcov, Phys. Lett. **B434**, 321 (1998), (E) *ibid.* **B444**, 584 (1998); see also: Nucl. Phys. (Proc. Supp.) **B77**, 93 (1999) and hep-ph/9811205.
155. M.V. Chizhov, M. Maris, and S.T. Petcov, hep-ph/9810501.
156. E.Kh. Akhmedov *et al.*, Nucl. Phys. **B542**, 3 (1999).
157. S.T. Petcov, Phys. Lett. **B214**, 259 (1988).
158. J. Hosaka *et al.*, [Super-Kamiokande Collab.], Phys. Rev. **D74**, 032002 (2006).
159. I. Mocioiu and R. Shrock, Phys. Rev. **D62**, 053017 (2000).
160. E.Kh. Akhmedov, Nucl. Phys. **B538**, 25 (1999).

161. M.V. Chizhov and S.T. Petcov, Phys. Rev. Lett. **83**, 1096 (1999) and Phys. Rev. Lett. **85**, 3979 (2000); Phys. Rev. **D63**, 073003 (2001).
162. J. Bernab  , S. Palomares-Ruiz, and S.T. Petcov, Nucl. Phys. **B669**, 255 (2003); S. Palomares-Ruiz and S.T. Petcov, Nucl. Phys. **B712**, 392 (2005); S.T. Petcov and T. Schwetz, Nucl. Phys. **B740**, 1 (2006); R. Gandhi *et al.*, Phys. Rev. **D76**, 073012 (2007); E.Kh. Akhmedov, M. Maltoni, and A.Yu. Smirnov, JHEP **0705**, 077 (2007).
163. The mantle-core enhancement maxima, *e.g.*, in  $P_m^{2\nu}(\nu_\mu \rightarrow \nu_\mu)$ , appeared in some of the early numerical calculations, but with incorrect interpretation (see, *e.g.*, the articles quoted in Ref. 152).
164. M. Honda *et al.*, Phys. Rev. **D52**, 4985 (1995); V. Agraval *et al.*, Phys. Rev. **D53**, 1314 (1996); G. Fiorentini *et al.*, Phys. Lett. **B510**, 173 (2001).
165. K. Abe *et al.*, (Letter of intent: Hyper-Kamiokande experiment), [arXiv:1109.3262](https://arxiv.org/abs/1109.3262).
166. O.L.G. Peres and A.Y. Smirnov, Phys. Lett. **B456**, 204 (1999), and Nucl. Phys. Proc. Suppl. **110**, 355(2002).
167. M. Freund, Phys. Rev. **D64**, 053003 (2001).
168. M.C. Gonzalez-Garcia and Y. Nir, Rev. Mod. Phys. **75**, 345 (2003); S.M. Bilenky, W. Grimus, and C. Giunti, Prog. in Part. Nucl. Phys. **43**, 1 (1999).
169. J.N. Bahcall, *Neutrino Astrophysics*, Cambridge University Press, Cambridge, 1989; J.N. Bahcall and M. Pinsonneault, Phys. Rev. Lett. **92**, 121301 (2004).
170. A. Messiah, *Proc. of the VIth Moriond Workshop* (eds. O. Fackler, J. Tran Thanh Van, Fronti  res, Gif-sur-Yvette, 1986), p. 373.
171. S.J. Parke, Phys. Rev. Lett. **57**, 1275 (1986).
172. S.T. Petcov, Phys. Lett. **B200**, 373 (1988).
173. P.I. Krastev and S.T. Petcov, Phys. Lett. **B207**, 64 (1988); M. Bruggen, W.C. Haxton, and Y.-Z. Quian, Phys. Rev. **D51**, 4028 (1995).
174. T. Kaneko, Prog. Theor. Phys. **78**, 532 (1987); S. Toshev, Phys. Lett. **B196**, 170 (1987); M. Ito, T. Kaneko, and M. Nakagawa, Prog. Theor. Phys. **79**, 13 (1988), (E) *ibid.*, **79**, 555 (1988).
175. S.T. Petcov, Phys. Lett. **B406**, 355 (1997).
176. C. Cohen-Tannoudji, B. Diu, and F. Laloe, *Quantum Mechanics*, Vol. 1 (Hermann, Paris, and John Wiley & Sons, New York, 1977).
177. S.T. Petcov, Phys. Lett. **B214**, 139 (1988); E. Lisi *et al.*, Phys. Rev. **D63**, 093002 (2000); A. Friedland, Phys. Rev. **D64**, 013008 (2001).
178. S.T. Petcov and J. Rich, Phys. Lett. **B224**, 401 (1989).
179. An expression for the “jump” probability  $P'$  for  $N_e$  varying linearly along the neutrino path was derived in W.C. Haxton, Phys. Rev. Lett. **57**, 1271 (1986) and in Ref. 171 on the basis of the old Landau-Zener result: L.D. Landau, Phys. Z. USSR **1**, 426 (1932), C. Zener, Proc. R. Soc. A **137**, 696 (1932). An analytic description of the solar  $\nu_e$  transitions based on the Landau-Zener jump probability was proposed in Ref. 171 and in W.C. Haxton, Phys. Rev. **D35**, 2352 (1987). The precision limitations of this description, which is less accurate than that based on the exponential density approximation, were discussed in S.T. Petcov, Phys. Lett. **B191**, 299 (1987) and in Ref. 173.
180. A. de Gouvea, A. Friedland, and H. Murayama, JHEP **0103**, 009 (2001).
181. C.-S. Lim, Report BNL 52079, 1987; S.P. Mikheev and A.Y. Smirnov, Phys. Lett. **B200**, 560 (1988).
182. G.L. Fogli *et al.*, Phys. Lett. **B583**, 149 (2004).
183. S.P. Mikheyev and A.Yu. Smirnov, *Proc. of the VIth Moriond Workshop* (eds. O. Fackler, J. Tran Thanh Van, Fronti  res, Gif-sur-Yvette, 1986), p. 355.
184. M. Cribier *et al.*, Phys. Lett. **B182**, 89 (1986); J. Bouchez *et al.*, Z. Phys. **C32**, 499(1986).
185. E. Lisi and D. Montanino, Phys. Rev. **D56**, 1792 (1997); Q.Y. Liu, M. Maris and S.T. Petcov, Phys. Rev. **D56**, 5991 (1997); M. Maris and S.T. Petcov, Phys. Rev. **D56**, 7444 (1997); J.N. Bahcall and P.I. Krastev, Phys. Rev. **C56**, 2839 (1997); J.N. Bahcall, P.I. Krastev and A.Y. Smirnov, Phys. Rev. **D60**, 093001 (1999); M. Maris and S.T. Petcov, Phys. Rev. **D62**, 093006 (2000); J.N. Bahcall, P.I. Krastev and A.Y. Smirnov, Phys. Rev. **D62**, 093004 (2000); M.C. Gonzalez-Garcia, C. Pena-Garay and A.Y. Smirnov, Phys. Rev. **D63**, 113004 (2001); P.I. Krastev and A.Y. Smirnov, Phys. Rev. **D65**, 073022 (2002).
186. A. Bandyopadhyay *et al.*, Phys. Lett. **B583**, 134 (2004); M. Blennow, T. Ohlsson and H. Snellman, Phys. Rev. **D69**, 073006 (2004); E.K. Akhmedov, M.A. Tortola and J.W.F. Valle, JHEP **0405**, 057 (2004).
187. K.S. Hirata *et al.*, [Kamiokande Collab.], Phys. Rev. Lett. **63**, 16 (1989).
188. Y. Fukuda *et al.*, [Super-Kamiokande Collab.], Phys. Rev. Lett. **81**, 1158 (1998).
189. J. Hosaka *et al.*, [Super-Kamiokande Collab.], Phys. Rev. **D73**, 112001 (2006).
190. J.P. Cravens *et al.*, [Super-Kamiokande Collab.], Phys. Rev. **D78**, 032002 (2008).
191. K. Abe *et al.*, [Super-Kamiokande Collab.], Phys. Rev. **D83**, 052010 (2011).
192. B. Aharmim *et al.*, [SNO Collab.], Phys. Rev. **C72**, 055502 (2005).
193. B. Aharmim *et al.*, [SNO Collab.], Phys. Rev. Lett. **101**, 111301 (2008); Phys. Rev. **C87**, 015502 (2013).
194. G. Bellini *et al.*, [Borexino Collab.], Phys. Rev. Lett. **107**, 141302 (2011).
195. G. Bellini *et al.*, [Borexino Collab.], Phys. Rev. Lett. **108**, 051302 (2012).
196. G. Bellini *et al.*, [Borexino Collab.], Nature **512**, 383 (2014).
197. G. Bellini *et al.*, [Borexino Collab.], Phys. Rev. **D82**, 033006 (2010).
198. G. Bellini *et al.*, [Borexino Collab.], Phys. Rev. **D89**, 112007 (2014).
199. A. Gando *et al.*, [KamLAND Collab.], Phys. Rev. **C92**, 055808 (2015).
200. S. Abe *et al.*, [KamLAND Collab.], Phys. Rev. **C84**, 035804 (2011).
201. B. Pontecorvo, Chalk River Lab. report PD-205, 1946.
202. P. Anselmann *et al.*, [GALLEX Collab.], Phys. Lett. **B342**, 440 (1995); W. Hampel *et al.*, [GALLEX Collab.], Phys. Lett. **B420**, 114 (1998).
203. J.N. Abdurashitov *et al.*, [SAGE Collab.], Phys. Rev. Lett. **77**, 4708 (1996); Phys. Rev. **C59**, 2246 (1999).
204. K.S. Hirata *et al.*, [Kamiokande Collab.], Phys. Lett. **B205**, 416 (1988) and Phys. Lett. **B280**, 146 (1992).
205. Y. Fukuda *et al.*, [Kamiokande Collab.], Phys. Lett. **B335**, 237 (1994).
206. D. Casper *et al.*, [IMB Collab.], Phys. Rev. Lett. **66**, 2561 (1991).
207. K. Daum *et al.*, [Frejus Collab.], Z. Phys. **C66**, 417 (1995).
208. W.W.M. Allison *et al.*, [Soudan 2 Collab.], Phys. Lett. **B391**, 491 (1997).
209. M. Ambrosio *et al.*, [MACRO Collab.], Phys. Lett. **B434**, 451 (1998).
210. P. Adamson *et al.*, [MINOS Collab.], Phys. Rev. **D86**, 052007 (2012).
211. M. Ageron *et al.*, [ANTARES Collab.], Nucl. Instrum. Methods **A656**, 11 (2011).
212. M.G. Aartsen *et al.*, [IceCube Collab.], Astropart. Phys. **35**, 615 (2012).
213. S. Adri  n-Mart  nez *et al.*, [ANTARES Collab.], Phys. Lett. **B714**, 224 (2012).
214. M.G. Aartsen *et al.*, [IceCube Collab.], Phys. Rev. **D91**, 072004 (2015).
215. M.G. Aartsen *et al.*, [IceCube Collab.], Phys. Rev. Lett. **111**, 081801 (2013).
216. G. Tzanakos *et al.*, [MINOS+ Collab.], FERMI LAB-PROPOSAL-1016 (2011).
217. See, <http://www.dunescience.org/>.
218. K. Abe *et al.*, Prog. Theor. Exp. Phys. **053C02** (2015).

219. S. Ahmed *et al.*, [INO Collab.], [arXiv:1505.07380](#).
220. M. G. Aartsen *et al.*, [IceCube-PINGU Collab.], [arXiv:1401.2046](#).
221. U.F. Katz (for the KM3NeT Collaboration), [arXiv:1402.1022](#).
222. S. Adrian-Martinez *et al.*, [KM3Net Collab.], [arXiv:1601.07459](#).
223. S. Wojcicki, *Proc. of the 1997 SLAC Summer Institute*.
224. M. Antonello *et al.*, [ICARUS Collab.], *Eur. Phys. J. C* **73**, 2345 (2013); *Eur. Phys. J. C* **73**, 2599 (2013).
225. A. Aguilar *et al.*, [LSND Collab.], *Phys. Rev. D* **64**, 112007 (2001).
226. B. Armbruster *et al.*, [Karmen Collab.], *Phys. Rev. D* **65**, 112001 (2002).
227. A.A. Aguilar-Arevalo *et al.*, [MiniBooNE Collab.], *Phys. Rev. Lett.* **98**, 231801 (2007); *Phys. Rev. Lett.* **102**, 101802 (2009).
228. A.A. Aguilar-Arevalo *et al.*, [MiniBooNE Collab.], *Phys. Rev. Lett.* **105**, 181801 (2010); *Phys. Rev. Lett.* **110**, 161801 (2013).
229. For the Fermilab short-Baseline Neutrino (SBN) Program, we refer to J. Asaadi, Talk at the WIN2015 International workshop.
230. F. Boem and P. Vogel, *Physics of Massive Neutrinos*, Cambridge Univ. Press, 1987.
231. F. Boehm *et al.*, *Phys. Rev. D* **64**, 112001 (2001).
232. Y.-F. Li *Int. J. Mod. Phys.: Conf. Ser.* **31**, 1460300 (2014).
233. D. Davis, Jr., D.S. Harmer, and K.C. Hoffman, *Phys. Rev. Lett.* **20**, 1205 (1968).
234. A.I. Abazov *et al.*, [SAGE Collab.], *Phys. Rev. Lett.* **67**, 3332 (1991).
235. Note that after publication of Ref. 8, the SAGE results including an extended observation period are reported by V.N. Gavrin in *Phys.-Usp.* **54**, 941 (2011). Also, F. Kaether *et al.*, *Phys. Lett. B* **685**, 47 (2010) reanalyzed a complete set of the GALLEX data with a method providing a better background reduction than that adopted in Ref. 10. The resulting GALLEX and GNO+GALLEX capture rates are slightly different from, but consistent with, those reported in Table 14.5..
236. Y. Fukuda *et al.*, [Super-Kamiokande Collab.], *Phys. Rev. Lett.* **86**, 5651 (2001).
237. Y. Fukuda *et al.*, [Super-Kamiokande Collab.], *Phys. Lett. B* **539**, 179 (2002).
238. G. L. Fogli *et al.*, *Phys. Rev. D* **67**, 073002 (2003); M. Maltoni, T. Schwetz, and J.W. Valle, *Phys. Rev. D* **67**, 093003 (2003); A. Bandyopadhyay *et al.*, *Phys. Lett. B* **559**, 121 (2003); J.N. Bahcall, M.C. Gonzalez-Garcia, and C. Peña-Garay, *JHEP* **0302**, 009 (2003); P.C. de Holanda and A.Y. Smirnov, *JCAP* **0302**, 001 (2003).
239. S. Abe *et al.*, [KamLAND Collab.], *Phys. Rev. Lett.* **100**, 221803 (2008).
240. A. Gando *et al.*, [KamLAND Collab.], *Phys. Rev. D* **83**, 052002 (2011).
241. B. Aharmim *et al.*, [SNO Collab.], *Phys. Rev. C* **81**, 055504 (2010).
242. B. Aharmim *et al.*, [SNO Collab.], *Phys. Rev. C* **88**, 025501 (2013).
243. G. Bellini *et al.*, [Borexino Collab.], *Phys. Lett. B* **707**, 22 (2012).
244. A. Renshaw *et al.*, [Super-Kamiokande Collab.], *Phys. Rev. Lett.* **112**, 091805 (2014).
245. S. T. Petcov, *Nucl. Phys. B* **892**, 400 (2015); C. Hagedorn, A. Meroni, and E. Molinaro, *Nucl. Phys. B* **891**, 499 (2015); I. Girardi, S.T. Petcov, and A.V. Titov, *Nucl. Phys. B* **894**, 733 (2015), and *Eur. Phys. J. C* **75**, 345 (2015); P. Ballett, S. Pascoli, and J. Turner, [arXiv:1503.07543](#); C.C. Li and G.J. Ding, [arXiv:1503.03711](#); I. Girardi *et al.*, [arXiv:1509.02502](#).
246. V. Barger *et al.*, *Phys. Rev. Lett.* **82**, 2640 (1999).
247. E. Lisi *et al.*, *Phys. Rev. Lett.* **85**, 1166 (2000).
248. W.W.M. Allison *et al.*, [Soudan 2 Collab.], *Phys. Rev. D* **68**, 113004 (2003).
249. K. Abe *et al.*, [T2K Collab.], *Phys. Rev. Lett.* **112**, 181801 (2014).
250. P. Adamson *et al.*, [MINOS Collab.], *Phys. Rev. Lett.* **110**, 251801 (2013).
251. A. Himmel (for the Super-Kamiokande Collab.), *AIP Conf. Proc.* **1604**, 345 (2014); [arXiv:1310.6677](#).
252. P. Adamson *et al.*, [MINOS Collab.], *Phys. Rev. Lett.* **112**, 191801 (2014).
253. P. Adamson *et al.*, [MINOS Collab.], *Phys. Rev. Lett.* **107**, 021801 (2011); *Phys. Rev. Lett.* **108**, 191801 (2012).
254. K. Abe *et al.*, [Super-Kamiokande Collab.], *Phys. Rev. Lett.* **97**, 171801 (2006).
255. N. Agafonova *et al.*, [OPERA Collab.], [arXiv:1507.01417](#).
256. J. Beringer *et al.*, *Phys. Rev. D* **86**, 010001 (2012).
257. G. Mention *et al.*, *Phys. Rev. D* **83**, 073006 (2011).
258. A.C. Hayes *et al.*, *Phys. Rev. Lett.* **112**, 202501 (2014).
259. P. Astier *et al.*, [NOMAD Collab.], *Phys. Lett. B* **570**, 19 (2003).
260. N. Agafanova *et al.*, [OPERA Collab.], *JHEP* **1307**, 004 (2013); *JHEP* **1307**, 085 (2013).
261. F. Dydak *et al.*, [CDHSW Collab.], *Phys. Lett. B* **134**, 281 (1984).
262. J. Kopp *et al.*, *JHEP* **1305**, 050 (2013).
263. C. Giunti *et al.*, *Phys. Rev. D* **88**, 073008 (2013).
264. F.P. An *et al.*, [Daya Bay Collab.], *Phys. Rev. Lett.* **113**, 141802 (2014).
265. M. Archidiacono *et al.*, *Phys. Rev. D* **86**, 065028 (2012).
266. S. Gariazzo *et al.*, [arXiv:1507.08204](#).

## 15. QUARK MODEL

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### 15.1. Quantum numbers of the quarks

Quantum chromodynamics (QCD) is the theory of the strong interactions. QCD is a quantum field theory and its constituents are a set of fermions, the quarks, and gauge bosons, the gluons. Strongly interacting particles, the hadrons, are bound states of quark and gluon fields. As gluons carry no intrinsic quantum numbers beyond color charge, and because color is believed to be permanently confined, most of the quantum numbers of strongly interacting particles are given by the quantum numbers of their constituent quarks and antiquarks. The description of hadronic properties which strongly emphasizes the role of the minimum-quark-content part of the wave function of a hadron is generically called the quark model. It exists on many levels: from the simple, almost dynamics-free picture of strongly interacting particles as bound states of quarks and antiquarks, to more detailed descriptions of dynamics, either through models or directly from QCD itself. The different sections of this review survey the many approaches to the spectroscopy of strongly interacting particles which fall under the umbrella of the quark model.

**Table 15.1:** Additive quantum numbers of the quarks.

	$d$	$u$	$s$	$c$	$b$	$t$
Q – electric charge	$-\frac{1}{3}$	$+\frac{2}{3}$	$-\frac{1}{3}$	$+\frac{2}{3}$	$-\frac{1}{3}$	$+\frac{2}{3}$
I – isospin	$\frac{1}{2}$	$\frac{1}{2}$	0	0	0	0
$I_z$ – isospin $z$ -component	$-\frac{1}{2}$	$+\frac{1}{2}$	0	0	0	0
S – strangeness	0	0	-1	0	0	0
C – charm	0	0	0	+1	0	0
B – bottomness	0	0	0	0	-1	0
T – topness	0	0	0	0	0	+1

Quarks are strongly interacting fermions with spin 1/2 and, by convention, positive parity. Antiquarks have negative parity. Quarks have the additive baryon number 1/3, antiquarks -1/3. Table 15.1 gives the other additive quantum numbers (flavors) for the three generations of quarks. They are related to the charge Q (in units of the elementary charge  $e$ ) through the generalized Gell-Mann-Nishijima formula

$$Q = I_z + \frac{B + S + C + B + T}{2}, \quad (15.1)$$

where  $B$  is the baryon number. The convention is that the *flavor* of a quark ( $I_z$ , S, C, B, or T) has the same sign as its *charge* Q. With this convention, any flavor carried by a charged meson has the same sign as its charge, *e.g.*, the strangeness of the  $K^+$  is +1, the bottomness of the  $B^+$  is +1, and the charm and strangeness of the  $D_s^-$  are each -1. Antiquarks have the opposite flavor signs. The hypercharge is defined as

$$Y = B + S - \frac{C - B + T}{3}.$$

Thus  $Y$  is equal to  $\frac{1}{3}$  for the  $u$  and  $d$  quarks,  $-\frac{2}{3}$  for the  $s$  quark, and 0 for all other quarks.

### 15.2. Mesons

Mesons have baryon number  $B = 0$ . In the quark model, they are  $q\bar{q}'$  bound states of quarks  $q$  and antiquarks  $\bar{q}'$  (the flavors of  $q$  and  $q'$  may be different). If the orbital angular momentum of the  $q\bar{q}'$  state is  $\ell$ , then the parity  $P$  is  $(-1)^{\ell+1}$ . The meson spin  $J$  is given by the usual relation  $|\ell - s| \leq J \leq |\ell + s|$ , where  $s$  is 0 (antiparallel quark spins) or 1 (parallel quark spins). The charge conjugation, or  $C$ -parity  $C = (-1)^{\ell+s}$ , is defined only for the  $q\bar{q}$  states made of quarks and their own antiquarks. The  $C$ -parity can be generalized to the  $G$ -parity  $G = (-1)^{I+\ell+s}$  for mesons made of quarks and their own antiquarks (isospin  $I_z = 0$ ), and for the charged  $u\bar{d}$  and  $d\bar{u}$  states (isospin  $I = 1$ ).

The mesons are classified in  $J^{PC}$  multiplets. The  $\ell = 0$  states are the pseudoscalars ( $0^{-+}$ ) and the vectors ( $1^{--}$ ). The orbital excitations  $\ell = 1$  are the scalars ( $0^{++}$ ), the axial vectors ( $1^{++}$ ) and ( $1^{+-}$ ), and the tensors ( $2^{++}$ ). Assignments for many of the known mesons are given in Tables 15.2 and 15.3. Radial excitations are denoted by the principal quantum number  $n$ . The very short lifetime of the  $t$  quark makes it likely that bound-state hadrons containing  $t$  quarks and/or antiquarks do not exist.

States in the natural spin-parity series  $P = (-1)^J$  must, according to the above, have  $s = 1$  and hence,  $CP = +1$ . Thus, mesons with natural spin-parity and  $CP = -1$  ( $0^{+-}$ ,  $1^{-+}$ ,  $2^{+-}$ ,  $3^{-+}$ , *etc.*) are forbidden in the  $q\bar{q}'$  model. The  $J^{PC} = 0^{--}$  state is forbidden as well. Mesons with such *exotic* quantum numbers may exist, but would lie outside the  $q\bar{q}'$  model (see section below on exotic mesons).

Following SU(3), the nine possible  $q\bar{q}'$  combinations containing the light  $u$ ,  $d$ , and  $s$  quarks are grouped into an octet and a singlet of light quark mesons:

$$\mathbf{3} \otimes \bar{\mathbf{3}} = \mathbf{8} \oplus \mathbf{1}. \quad (15.2)$$

A fourth quark such as charm  $c$  can be included by extending SU(3) to SU(4). However, SU(4) is badly broken owing to the much heavier  $c$  quark. Nevertheless, in an SU(4) classification, the sixteen mesons are grouped into a 15-plet and a singlet:

$$\mathbf{4} \otimes \bar{\mathbf{4}} = \mathbf{15} \oplus \mathbf{1}. \quad (15.3)$$

The *weight diagrams* for the ground-state pseudoscalar ( $0^{-+}$ ) and vector ( $1^{--}$ ) mesons are depicted in Fig. 15.1. The light quark mesons are members of nonets building the middle plane in Fig. 15.1(a) and (b).

Isoscalar states with the same  $J^{PC}$  will mix, but mixing between the two light quark isoscalar mesons, and the much heavier charmonium or bottomonium states, are generally assumed to be negligible. In the following, we shall use the generic names  $a$  for the  $I = 1$ ,  $K$  for the  $I = 1/2$ , and  $f$  and  $f'$  for the  $I = 0$  members of the light quark nonets. Thus, the physical isoscalars are mixtures of the SU(3) wave function  $\psi_8$  and  $\psi_1$ :

$$f' = \psi_8 \cos \theta - \psi_1 \sin \theta, \quad (15.4)$$

$$f = \psi_8 \sin \theta + \psi_1 \cos \theta, \quad (15.5)$$

where  $\theta$  is the nonet mixing angle and

$$\psi_8 = \frac{1}{\sqrt{6}}(u\bar{u} + d\bar{d} - 2s\bar{s}), \quad (15.6)$$

$$\psi_1 = \frac{1}{\sqrt{3}}(u\bar{u} + d\bar{d} + s\bar{s}). \quad (15.7)$$

These mixing relations are often rewritten to exhibit the  $u\bar{u} + d\bar{d}$  and  $s\bar{s}$  components which decouple for the “ideal” mixing angle  $\theta_i$ , such that  $\tan \theta_i = 1/\sqrt{2}$  (or  $\theta_i = 35.3^\circ$ ). Defining  $\alpha = \theta + 54.7^\circ$ , one obtains the physical isoscalar in the flavor basis

$$f' = \frac{1}{\sqrt{2}}(u\bar{u} + d\bar{d}) \cos \alpha - s\bar{s} \sin \alpha, \quad (15.8)$$

**Table 15.2:** Suggested  $q\bar{q}$  quark-model assignments for some of the observed light mesons. Mesons in bold face are included in the Meson Summary Table. The wave functions  $f$  and  $f'$  are given in the text. The singlet-octet mixing angles from the quadratic and linear mass formulae are also given for the well established nonets. The classification of the  $0^{++}$  mesons is tentative: the light scalars  $a_0(980)$ ,  $f_0(980)$ ,  $f_0(500)$  and  $K_0^*(800)$  are often considered to be meson-meson resonances or four-quark states, and are omitted from the table. The isoscalar  $0^{++}$  mesons are expected to mix. In particular, the  $f_0(1710)$  mixes with the  $f_0(1500)$  and the  $f_0(1370)$ . The  $a_0(1450)$  is not firmly established. See the “Note on Non- $q\bar{q}$  mesons” and the “Note on Scalar Mesons” in the Meson Listings for details and alternative schemes. In the  $1^{++}$  nonet the isoscalar slot is disputed by the  $f_1(1510)$ . The isoscalar assignments in the  $2^1S_0$  ( $0^{++}$ ) nonet are also tentative. See the “Note on The Pseudoscalar and Pseudovector Mesons in the 1400 MeV Region” in the Meson Listings.

$n^{2s+1}\ell_J$	$J^{PC}$	$l = 1$ $u\bar{d}, \bar{u}d, \frac{1}{\sqrt{2}}(d\bar{d} - u\bar{u})$	$l = \frac{1}{2}$ $u\bar{s}, d\bar{s}, \bar{d}s, -\bar{u}s$	$l = 0$ $f'$	$l = 0$ $f$	$\theta_{\text{quad}}$ [°]	$\theta_{\text{lin}}$ [°]
$1^1S_0$	$0^{-+}$	$\pi$	$K$	$\eta$	$\eta'(958)$	-11.4	-24.5
$1^3S_1$	$1^{--}$	$\rho(770)$	$K^*(892)$	$\phi(1020)$	$\omega(782)$	39.1	36.4
$1^1P_1$	$1^{+-}$	$b_1(1235)$	$K_{1B}^\dagger$	$h_1(1380)$	$h_1(1170)$		
$1^3P_0$	$0^{++}$	$a_0(1450)$	$K_0^*(1430)$	$f_0(1710)$	$f_0(1370)$		
$1^3P_1$	$1^{++}$	$a_1(1260)$	$K_{1A}^\dagger$	$f_1(1420)$	$f_1(1285)$		
$1^3P_2$	$2^{++}$	$a_2(1320)$	$K_2^*(1430)$	$f_2'(1525)$	$f_2(1270)$	32.1	30.5
$1^1D_2$	$2^{-+}$	$\pi_2(1670)$	$K_2(1770)^\dagger$	$\eta_2(1870)$	$\eta_2(1645)$		
$1^3D_1$	$1^{--}$	$\rho(1700)$	$K^*(1680)$		$\omega(1650)$		
$1^3D_2$	$2^{--}$		$K_2(1820)$				
$1^3D_3$	$3^{--}$	$\rho_3(1690)$	$K_3^*(1780)$	$\phi_3(1850)$	$\omega_3(1670)$	31.8	30.8
$1^3F_4$	$4^{++}$	$a_4(2040)$	$K_4^*(2045)$		$f_4(2050)$		
$1^3G_5$	$5^{--}$	$\rho_5(2350)$	$K_5^*(2380)$				
$1^3H_6$	$6^{++}$	$a_6(2450)$			$f_6(2510)$		
$2^1S_0$	$0^{-+}$	$\pi(1300)$	$K(1460)$	$\eta(1475)$	$\eta(1295)$		
$2^3S_1$	$1^{--}$	$\rho(1450)$	$K^*(1410)$	$\phi(1680)$	$\omega(1420)$		

$^\dagger$  The  $1^{++}$  and  $2^{-+}$  isospin  $\frac{1}{2}$  states mix. In particular, the  $K_{1A}$  and  $K_{1B}$  are nearly equal ( $45^\circ$ ) mixtures of the  $K_1(1270)$  and  $K_1(1400)$ . The physical vector mesons listed under  $1^3D_1$  and  $2^3S_1$  may be mixtures of  $1^3D_1$  and  $2^3S_1$ .

**Table 15.3:**  $q\bar{q}$  quark-model assignments for the observed heavy mesons with established  $J^{PC}$ . Mesons in bold face are included in the Meson Summary Table.

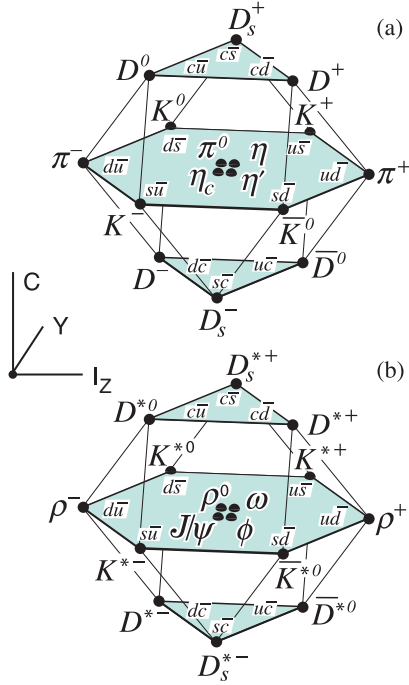
$n^{2s+1}\ell_J$	$J^{PC}$	$l = 0$ $c\bar{c}$	$l = 0$ $b\bar{b}$	$l = \frac{1}{2}$ $c\bar{u}, c\bar{d}; \bar{c}u, \bar{c}d$	$l = 0$ $c\bar{s}; \bar{c}s$	$l = \frac{1}{2}$ $b\bar{u}, b\bar{d}; \bar{b}u, \bar{b}d$	$l = 0$ $b\bar{s}; \bar{b}s$	$l = 0$ $b\bar{c}; \bar{b}c$
$1^1S_0$	$0^{-+}$	$\eta_c(1S)$	$\eta_b(1S)$	$D$	$D_s^\pm$	$B$	$B_s^0$	$B_c^\pm$
$1^3S_1$	$1^{--}$	$J/\psi(1S)$	$\Upsilon(1S)$	$D^*$	$D_s^{*\pm}$	$B^*$	$B_s^*$	
$1^1P_1$	$1^{+-}$	$h_c(1P)$	$h_b(1P)$	$D_1(2420)$	$D_{s1}(2536)^\pm$	$B_1(5721)$	$B_{s1}(5830)^0$	
$1^3P_0$	$0^{++}$	$\chi_{c0}(1P)$	$\chi_{b0}(1P)$	$D_0^*(2400)$	$D_{s0}^*(2317)^\pm$			
$1^3P_1$	$1^{++}$	$\chi_{c1}(1P)$	$\chi_{b1}(1P)$	$D_1(2430)$	$D_{s1}(2460)^\pm$			
$1^3P_2$	$2^{++}$	$\chi_{c2}(1P)$	$\chi_{b2}(1P)$	$D_2^*(2460)$	$D_{s2}^*(2573)^\pm$	$B_2^*(5747)$	$B_{s2}^*(5840)^0$	
$1^3D_1$	$1^{--}$	$\psi(3770)$			$D_{s1}^*(2860)^\pm$			
$1^3D_3$	$3^{--}$				$D_{s3}^*(2860)^\pm$			
$2^1S_0$	$0^{-+}$	$\eta_c(2S)$	$\eta_b(2S)$	$D(2550)$				
$2^3S_1$	$1^{--}$	$\psi(2S)$	$\Upsilon(2S)$		$D_{s1}^*(2700)^\pm$			
$2^1P_1$	$1^{+-}$		$h_b(2P)$					
$2^3P_{0,1,2}$	$0^{++}, 1^{++}, 2^{++}$	$\chi_{c0,2}(2P)$	$\chi_{b0,1,2}(2P)$					
$3^3P_{0,1,2}$	$0^{++}, 1^{++}, 2^{++}$		$\chi_b(3P)$					

$^\dagger$  The masses of these states are considerably smaller than most theoretical predictions. They have also been considered as four-quark states.

$^\ddagger$  These states are mixtures of the  $1^3D_1$  and  $2^3S_1$  states.

The open flavor states in the  $1^{+-}$  and  $1^{++}$  rows are mixtures of the  $1^{++}$  states.





**Figure 15.1:** SU(4) weight diagram showing the 16-plets for the pseudoscalar (a) and vector mesons (b) made of the  $u$ ,  $d$ ,  $s$ , and  $c$  quarks as a function of isospin  $I_z$ , charm  $C$ , and hypercharge  $Y = B + S - \frac{C}{3}$ . The nonets of light mesons occupy the central planes to which the  $c\bar{c}$  states have been added.

and its orthogonal partner  $f$  (replace  $\alpha$  by  $\alpha - 90^\circ$ ). Thus for ideal mixing ( $\alpha_i = 90^\circ$ ), the  $f'$  becomes pure  $s\bar{s}$  and the  $f$  pure  $u\bar{u} + d\bar{d}$ . The mixing angle  $\theta$  can be derived by diagonalizing the mass matrix

$$\begin{pmatrix} m_8 & m_{81} \\ m_{18} & m_1 \end{pmatrix}$$

The mass eigenvalues are  $m_{f'}$  and  $m_f$ . The mixing angle is given by

$$\tan \theta = \frac{m_8 - m_{f'}}{m_{81}}.$$

Calculating  $m_8$  and  $m_{81}$  from the wave functions Eq. (15.6) and Eq. (15.7), and expressing the quark masses as a function of the  $l = 1/2$  and  $l = 1$  meson masses, one obtains

$$\tan \theta = \frac{4m_K - m_a - 3m_{f'}}{2\sqrt{2}(m_a - m_K)}, \quad (15.9)$$

which also determines the sign of  $\theta$ . Alternatively, one can express the mixing angle as a function of all nonet masses. The octet mass is given by

$$m_8 = m_{f'} \cos^2 \theta + m_f \sin^2 \theta$$

whence

$$\tan^2 \theta = \frac{4m_K - m_a - 3m_{f'}}{-4m_K + m_a + 3m_f}. \quad (15.10)$$

Eliminating  $\theta$  from Eq. (15.9) and Eq. (15.10) leads to the sum rule [1]

$$(m_f + m_{f'})(4m_K - m_a) - 3m_f m_{f'} = 8m_K^2 - 8m_K m_a + 3m_a^2. \quad (15.11)$$

This relation is verified for the ground-state vector mesons. We identify the  $\phi(1020)$  with the  $f'$  and the  $\omega(783)$  with the  $f$ . Thus

$$\phi(1020) = \psi_8 \cos \theta_V - \psi_1 \sin \theta_V, \quad (15.12)$$

$$\omega(783) = \psi_8 \sin \theta_V + \psi_1 \cos \theta_V, \quad (15.13)$$

with the vector mixing angle  $\theta_V = 36.4^\circ$  from Eq. (15.10), very close to ideal mixing. Thus  $\phi(1020)$  is nearly pure  $s\bar{s}$ . For ideal mixing, Eq. (15.9) and Eq. (15.10) lead to the relations

$$m_K = \frac{m_f + m_{f'}}{2}, \quad m_a = m_f, \quad (15.14)$$

which are satisfied for the vector mesons.

The situation for the pseudoscalar and scalar mesons is not so clear cut, either theoretically or experimentally. For the pseudoscalars, the mixing angle is small. This can be understood qualitatively via gluon-line counting of the mixing process. The size of the mixing process between the nonstrange and strange mass bases scales as  $\alpha_s^2$ , not  $\alpha_s^3$ , because of two rather than three gluon exchange as it does for the vector mesons. It may also be that the lightest isoscalar pseudoscalars mix more strongly with excited states or with states of substantial non- $\bar{q}q$  content, as will be discussed below.

A variety of analysis methods lead to similar results: First, for these states, Eq. (15.11) is satisfied only approximately. Then Eq. (15.9) and Eq. (15.10) lead to somewhat different values for the mixing angle. Identifying the  $\eta$  with the  $f'$  one gets

$$\eta = \psi_8 \cos \theta_P - \psi_1 \sin \theta_P, \quad (15.15)$$

$$\eta' = \psi_8 \sin \theta_P + \psi_1 \cos \theta_P. \quad (15.16)$$

Following chiral perturbation theory, the meson masses in the mass formulae (Eq. (15.9) and Eq. (15.10)) might be replaced by their squares. Table 15.2 lists the mixing angle  $\theta_{\text{lin}}$  from Eq. (15.10) (using the neutral members of the nonets) and the corresponding  $\theta_{\text{quad}}$  obtained by replacing the meson masses by their squares throughout.

The pseudoscalar mixing angle  $\theta_P$  can also be measured by comparing the partial widths for radiative  $J/\psi$  decay into a vector and a pseudoscalar [2], radiative  $\phi(1020)$  decay into  $\eta$  and  $\eta'$  [3], or  $\bar{p}p$  annihilation at rest into a pair of vector and pseudoscalar or into two pseudoscalars [4,5]. One obtains a mixing angle between  $-10^\circ$  and  $-20^\circ$ . More recently, a lattice QCD simulation, Ref. 6, has successfully reproduced the masses of the  $\eta$  and  $\eta'$ , and as a byproduct find a mixing angle  $\theta_{\text{lin}} = -14.1(2.8)^\circ$ . We return to this point in Sec. 15.6.

The nonet mixing angles can be measured in  $\gamma\gamma$  collisions, *e.g.*, for the  $0^{-+}$ ,  $0^{++}$ , and  $2^{++}$  nonets. In the quark model, the amplitude for the coupling of neutral mesons to two photons is proportional to  $\sum_i Q_i^2$ , where  $Q_i$  is the charge of the  $i$ -th quark. The  $2\gamma$  partial width of an isoscalar meson with mass  $m$  is then given in terms of the mixing angle  $\alpha$  by

$$\Gamma_{2\gamma} = C(5 \cos \alpha - \sqrt{2} \sin \alpha)^2 m^3, \quad (15.17)$$

for  $f'$  and  $f$  ( $\alpha \rightarrow \alpha - 90^\circ$ ). The coupling  $C$  may depend on the meson mass. It is often assumed to be a constant in the nonet. For the isovector  $a$ , one then finds  $\Gamma_{2\gamma} = 9 C m^3$ . Thus the members of an ideally mixed nonet couple to  $2\gamma$  with partial widths in the ratios  $f' : a = 25 : 2 : 9$ . For tensor mesons, one finds from the ratios of the measured  $2\gamma$  partial widths for the  $f_2(1270)$  and  $f_2'(1525)$  mesons a mixing angle  $\alpha_T$  of  $(81 \pm 1)^\circ$ , or  $\theta_T = (27 \pm 1)^\circ$ , in accord with the linear mass formula. For the pseudoscalars, one finds from the ratios of partial widths  $\Gamma(\eta' \rightarrow 2\gamma)/\Gamma(\eta \rightarrow 2\gamma)$  a mixing angle  $\theta_P = (-18 \pm 2)^\circ$ , while the ratio  $\Gamma(\eta' \rightarrow 2\gamma)/\Gamma(\pi^0 \rightarrow 2\gamma)$  leads to  $\sim -24^\circ$ . SU(3) breaking effects for pseudoscalars are discussed in Ref. 7.

The partial width for the decay of a scalar or a tensor meson into a pair of pseudoscalar mesons is model-dependent. Following Ref. 8,

$$\Gamma = C \times \gamma^2 \times |F(q)|^2 \times q. \quad (15.18)$$

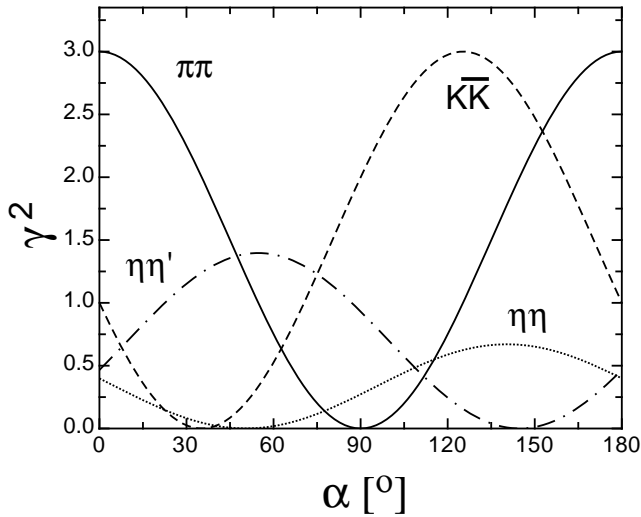
$C$  is a nonet constant,  $q$  the momentum of the decay products,  $F(q)$  a form factor, and  $\gamma^2$  the SU(3) coupling. The model-dependent form factor may be written as

$$|F(q)|^2 = q^{2\ell} \times \exp\left(-\frac{q^2}{8\beta^2}\right), \quad (15.19)$$

where  $\ell$  is the relative angular momentum between the decay products. The decay of a  $q\bar{q}$  meson into a pair of mesons involves the creation of a  $q\bar{q}$  pair from the vacuum, and SU(3) symmetry assumes that the matrix elements for the creation of  $s\bar{s}$ ,  $u\bar{u}$ , and  $d\bar{d}$  pairs are equal. The couplings  $\gamma^2$  are given in Table 15.4, and their dependence upon the mixing angle  $\alpha$  is shown in Fig. 15.2 for isoscalar decays. The generalization to unequal  $s\bar{s}$ ,  $u\bar{u}$ , and  $d\bar{d}$  couplings is given in Ref. 8. An excellent fit to the tensor meson decay widths is obtained assuming SU(3) symmetry, with  $\beta \simeq 0.5 \text{ GeV}/c$ ,  $\theta_V \simeq 26^\circ$  and  $\theta_P \simeq -17^\circ$  [8].

**Table 15.4:** SU(3) couplings  $\gamma^2$  for quarkonium decays as a function of nonet mixing angle  $\alpha$ , up to a common multiplicative factor  $C$  ( $\phi \equiv 54.7^\circ + \theta_P$ ).

Isospin	Decay channel	$\gamma^2$
0	$\pi\pi$	$3\cos^2\alpha$
	$K\bar{K}$	$(\cos\alpha - \sqrt{2}\sin\alpha)^2$
	$\eta\eta$	$(\cos\alpha\cos^2\phi - \sqrt{2}\sin\alpha\sin^2\phi)^2$
	$\eta\eta'$	$\frac{1}{2}\sin^2 2\phi (\cos\alpha + \sqrt{2}\sin\alpha)^2$
1	$\eta\pi$	$2\cos^2\phi$
	$\eta'\pi$	$2\sin^2\phi$
	$K\bar{K}$	1
$\frac{1}{2}$	$K\pi$	$\frac{3}{2}$
	$K\eta$	$(\sin\phi - \frac{\cos\phi}{\sqrt{2}})^2$
	$K\eta'$	$(\cos\phi + \frac{\sin\phi}{\sqrt{2}})^2$



**Figure 15.2:** SU(3) couplings as a function of mixing angle  $\alpha$  for isoscalar decays, up to a common multiplicative factor  $C$  and for  $\theta_P = -17.3^\circ$ .

### 15.3. Exotic mesons

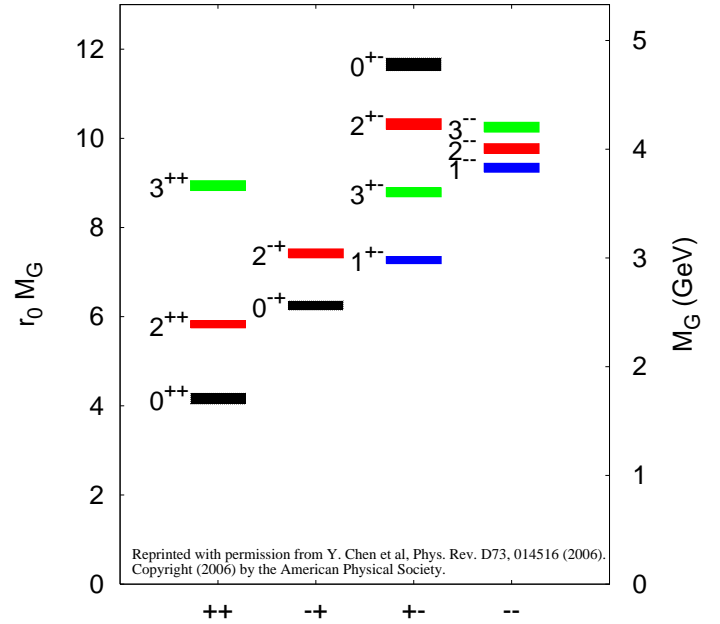
The existence of a light nonet composed of four quarks with masses below 1 GeV was suggested a long time ago [9]. Coupling two triplets of light quarks  $u$ ,  $d$ , and  $s$ , one obtains nine states, of which the six symmetric ( $uu$ ,  $dd$ ,  $ss$ ,  $ud + du$ ,  $us + su$ ,  $ds + sd$ ) form the six dimensional representation  $\mathbf{6}$ , while the three antisymmetric ( $ud - du$ ,  $us - su$ ,  $ds - sd$ ) form the three dimensional representation  $\bar{\mathbf{3}}$  of SU(3):

$$\mathbf{3} \otimes \mathbf{3} = \mathbf{6} \oplus \bar{\mathbf{3}}. \quad (15.20)$$

Combining with spin and color and requiring antisymmetry, one finds that the most deeply bound diquark (and hence the lightest) is the one in the  $\bar{\mathbf{3}}$  and spin singlet state. The combination of the diquark with an antiquark in the  $\mathbf{3}$  representation then gives a light nonet of four-quark scalar states. Letting the number of strange quarks determine the mass splitting, one obtains a mass inverted spectrum with a light isosinglet ( $u\bar{d}\bar{u}\bar{d}$ ), a medium heavy isodoublet (e.g.,  $u\bar{d}\bar{s}\bar{d}$ ) and a heavy isotriplet (e.g.,  $ds\bar{u}\bar{s}$ ) + isosinglet (e.g.,  $us\bar{u}\bar{s}$ ). It is then tempting to identify the lightest state with the  $f_0(500)$ , and the

heaviest states with the  $a_0(980)$ , and  $f_0(980)$ . Then the meson with strangeness  $K_0^*(800)$  would lie in-between.

QCD predicts the existence of extra isoscalar mesons. In the pure gauge theory they contain only gluons, and are called the glueballs. The ground state glueball is predicted by lattice gauge theories to be  $0^{++}$ , the first excited state  $2^{++}$ . Errors on the mass predictions are large. From Ref. 10 one obtains 1750 (50) (80) MeV for the mass of the lightest  $0^{++}$  glueball from quenched QCD. As an example for the glueball mass spectrum, we show in Fig. 15.3 a calculation from Ref. 11. A mass of 1710 MeV is predicted for the ground state, also with an error of about 100 MeV. Earlier work by other groups produced masses at 1650 MeV [12] and 1550 MeV [13] (see also [14]). The first excited state has a mass of about 2.4 GeV, and the lightest glueball with exotic quantum numbers ( $2^{+-}$ ) has a mass of about 4 GeV.



**Figure 15.3:** Predicted glueball mass spectrum from the lattice in quenched approximation (from Ref. 11)

These calculations are made in the so-called “quenched approximation” which neglects  $q\bar{q}$  loops. However, both glue and  $q\bar{q}$  states will couple to singlet scalar mesons. Therefore glueballs will mix with nearby  $q\bar{q}$  states of the same quantum numbers. For example, the two isoscalar  $0^{++}$  mesons around 1500 MeV will mix with the pure ground state glueball to generate the observed physical states  $f_0(1370)$ ,  $f_0(1500)$ , and  $f_0(1710)$  [8,15]. The first results from lattice calculations, which include these effects, indicate that the mass shifts are small. We return to a discussion of this point in Sec. 15.6.

The existence of three singlet scalar mesons around 1.5 GeV suggests additional degrees of freedom such as glue, since only two mesons are predicted in this mass range. The  $f_0(1500)$  [8,15] or, alternatively, the  $f_0(1710)$  [12], have been proposed as candidates for the scalar glueball, both states having considerable mixing also with the  $f_0(1370)$ . Other mixing schemes, in particular with the  $f_0(500)$  and the  $f_0(980)$ , have also been proposed [16]. Details can be found in the “Note on Non- $q\bar{q}$  Mesons” in the Meson Listings and in Ref. 17. See also the “Note on Scalar Mesons below 2 GeV”.

Mesons made of  $q\bar{q}$  pairs bound by excited gluons  $g$ , the hybrid states  $q\bar{q}g$ , are also predicted. They should lie in the 1.9 GeV mass region, according to gluon flux tube models [18]. Lattice QCD also predicts the lightest hybrid, an exotic  $1^{-+}$ , at a mass of 1.8 to 1.9 GeV [19]. However, the bag model predicts four nonets, among them an exotic  $1^{-+}$  around or above 1.4 GeV [20,21]. There are so far two candidates for exotic states with quantum numbers  $1^{-+}$ , the  $\pi_1(1400)$  and  $\pi_1(1600)$ , which could be hybrids or four-quark states (see the “Note on Non- $q\bar{q}$  Mesons” in the Meson Listings and in Ref. 17).

### 15.4. Baryons: $qqq$ states

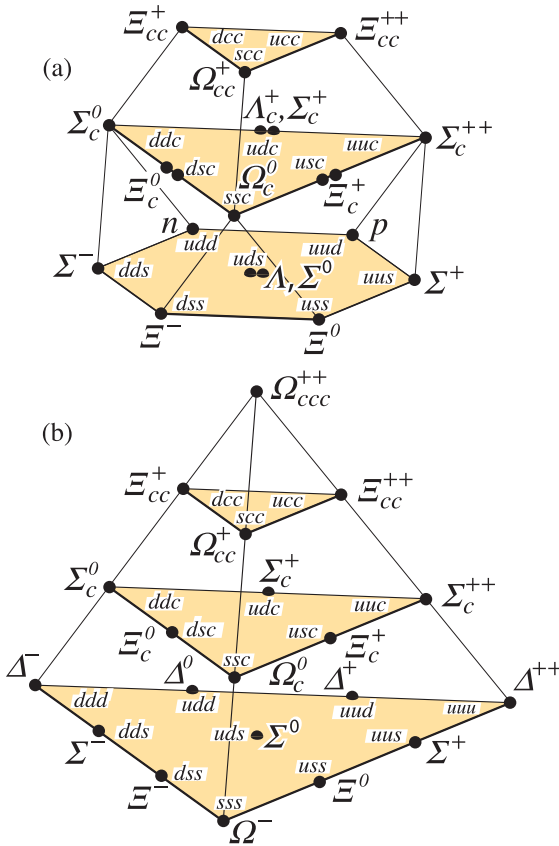
Baryons are fermions with baryon number  $\mathcal{B} = 1$ , *i.e.*, in the most general case, they are composed of three quarks plus any number of quark - antiquark pairs. So far all established baryons are 3-quark ( $qqq$ ) configurations (the LHCb collaboration has very recently announced observation of two charmed ‘pentaquark’ states of minimal quark content  $c\bar{c}uud$  at invariant masses close to 4.4 GeV [23]). The color part of their state functions is an SU(3) singlet, a completely antisymmetric state of the three colors. Since the quarks are fermions, the state function must be antisymmetric under interchange of any two equal-mass quarks (up and down quarks in the limit of isospin symmetry). Thus it can be written as

$$|qqq\rangle_A = |\text{color}\rangle_A \times |\text{space, spin, flavor}\rangle_S, \quad (15.21)$$

where the subscripts  $S$  and  $A$  indicate symmetry or antisymmetry under interchange of any two equal-mass quarks. Note the contrast with the state function for the three nucleons in  ${}^3\text{H}$  or  ${}^3\text{He}$ :

$$|NNN\rangle_A = |\text{space, spin, isospin}\rangle_A. \quad (15.22)$$

This difference has major implications for internal structure, magnetic moments, *etc.* (For a nice discussion, see Ref. 24)



**Figure 15.4:** SU(4) multiplets of baryons made of  $u$ ,  $d$ ,  $s$ , and  $c$  quarks. (a) The 20-plet with an SU(3) octet. (b) The 20-plet with an SU(3) decuplet.

The ‘ordinary’ baryons are made up of  $u$ ,  $d$ , and  $s$  quarks. The three flavors imply an approximate flavor SU(3), which requires that baryons made of these quarks belong to the multiplets on the right side of

$$\mathbf{3} \otimes \mathbf{3} \otimes \mathbf{3} = \mathbf{10}_S \oplus \mathbf{8}_M \oplus \mathbf{8}_M \oplus \mathbf{1}_A \quad (15.23)$$

(see Sec. 46, on ‘SU( $n$ ) Multiplets and Young Diagrams’). Here the subscripts indicate symmetric, mixed-symmetry, or antisymmetric states under interchange of any two quarks. The  $\mathbf{1}$  is a  $uds$  state ( $\Lambda_1$ ),

and the octet contains a similar state ( $\Lambda_8$ ). If these have the same spin and parity, they can mix. The mechanism is the same as for the mesons (see above). In the ground state multiplet, the SU(3) flavor singlet  $\Lambda_1$  is forbidden by Fermi statistics. Section 45, on ‘SU(3) Isoscalar Factors and Representation Matrices,’ shows how relative decay rates in, say,  $\mathbf{10} \rightarrow \mathbf{8} \otimes \mathbf{8}$  decays may be calculated.

The addition of the  $c$  quark to the light quarks extends the flavor symmetry to SU(4). However, due to the large mass of the  $c$  quark, this symmetry is much more strongly broken than the SU(3) of the three light quarks. Figures 15.4(a) and 15.4(b) show the SU(4) baryon multiplets that have as their bottom levels an SU(3) octet, such as the octet that includes the nucleon, or an SU(3) decuplet, such as the decuplet that includes the  $\Delta(1232)$ . All particles in a given SU(4) multiplet have the same spin and parity. The charmed baryons are discussed in more detail in the ‘Note on Charmed Baryons’ in the Particle Listings. The same multiplets as shown in 15.4 can be constructed when the  $c$  quark is replaced by the  $b$  quark, or they can be embedded in a larger SU(5) group that accounts for all baryons that can be constructed from the five quark flavors. The existence of baryons with  $t$ -quarks is very unlikely due to the short lifetime of the  $t$ -quark. The heavy quark baryons have recently gained a lot of interest. Their relatively narrow widths allow to isolate the states much easier than the light quark baryon resonances which require intricate partial wave analyses. The only problem on the experimental side are the small production cross sections, but the recent measurements at the  $e^+e^-$  colliding  $B$  factories, at the  $p\bar{p}$  Tevatron collider, and at LHCb at CERN have boosted this field. A recent summary is given in Ref. 25. A possible candidate for a doubly charmed baryon had been reported by the SELEX experiment [26,27] but could so far not be confirmed by other experiments, and quark model predictions for baryons with two heavy quarks are given in Ref. 28.

For the ‘ordinary’ baryons (no  $c$  or  $b$  quark), flavor and spin may be combined in an approximate flavor-spin SU(6), in which the six basic states are  $d \uparrow$ ,  $d \downarrow$ ,  $\dots$ ,  $s \downarrow$  ( $\uparrow, \downarrow$  = spin up, down). Then the baryons belong to the multiplets on the right side of

$$\mathbf{6} \otimes \mathbf{6} \otimes \mathbf{6} = \mathbf{56}_S \oplus \mathbf{70}_M \oplus \mathbf{70}_M \oplus \mathbf{20}_A. \quad (15.24)$$

These SU(6) multiplets decompose into flavor SU(3) multiplets as follows:

$$\mathbf{56} = \mathbf{4} \mathbf{10} \oplus \mathbf{2} \mathbf{8} \quad (15.25a)$$

$$\mathbf{70} = \mathbf{2} \mathbf{10} \oplus \mathbf{4} \mathbf{8} \oplus \mathbf{2} \mathbf{8} \oplus \mathbf{2} \mathbf{1} \quad (15.25b)$$

$$\mathbf{20} = \mathbf{2} \mathbf{8} \oplus \mathbf{4} \mathbf{1}, \quad (15.25c)$$

where the superscript  $(2S+1)$  gives the net spin  $S$  of the quarks for each particle in the SU(3) multiplet. The  $J^P = 1/2^+$  octet containing the nucleon and the  $J^P = 3/2^+$  decuplet containing the  $\Delta(1232)$  together make up the ‘ground-state’ 56-plet, in which the orbital angular momenta between the quark pairs are zero (so that the spatial part of the state function is trivially symmetric). The  $\mathbf{70}$  and  $\mathbf{20}$  require some excitation of the spatial part of the state function in order to make the overall state function symmetric. States with nonzero orbital angular momenta are classified in SU(6)  $\otimes$  O(3) supermultiplets.

It is useful to classify the baryons into bands that have the same number  $N$  of quanta of excitation. Each band consists of a number of supermultiplets, specified by  $(D, L_N^P)$ , where  $D$  is the dimensionality of the SU(6) representation,  $L$  is the total quark orbital angular momentum, and  $P$  is the total parity. Supermultiplets contained in bands up to  $N = 12$  are given in Ref. 29. The  $N = 0$  band, which contains the nucleon and  $\Delta(1232)$ , consists only of the  $(56, 0_0^+)$  supermultiplet. The  $N = 1$  band consists only of the  $(70, 1_1^-)$  multiplet and contains the negative-parity baryons with masses below about 1.9 GeV. The  $N = 2$  band contains five supermultiplets:  $(56, 0_2^+)$ ,  $(70, 0_2^+)$ ,  $(56, 2_2^+)$ ,  $(70, 2_2^+)$ , and  $(20, 1_2^+)$ .

The wave functions of the non-strange baryons in the harmonic oscillator basis are often labeled by  $|X^{2S+1} L_\pi J^P\rangle$ , where  $S, L, J, P$  are as above,  $X = N$  or  $\Delta$ , and  $\pi = S, M$  or  $A$  denotes the symmetry of the spatial wave function. The possible model states for the bands with  $N=0,1,2$  are given in Table 15.5. The assignment of experimentally observed states is only complete and well established

**Table 15.5:**  $N$  and  $\Delta$  states in the  $N=0,1,2$  harmonic oscillator bands.  $L^P$  denotes angular momentum and parity,  $S$  the three-quark spin and 'sym'=A,S,M the symmetry of the spatial wave function. Only dominant components indicated. Assignments in the  $N=2$  band are partly tentative.

$N$	sym	$L^P$	$S$	$N(I=1/2)$			$\Delta(I=3/2)$		
2	A	$1^+$	$1/2$	$1/2^+$	$3/2^+$				
2	M	$2^+$	$3/2$	$1/2^+$	$3/2^+$	$5/2^+$	$7/2^+$		
2	M	$2^+$	$1/2$		$3/2^+$	$5/2^+$		$3/2^+$	$5/2^+$
2	M	$0^+$	$3/2$		$3/2^+$				
2	M	$0^+$	$1/2$	$1/2^+$			$1/2^+$		
				$N(1710)$			$\Delta(1750)$		
2	S	$2^+$	$3/2$				$1/2^+$	$3/2^+$	$5/2^+$
							$\Delta(1910)$	$\Delta(1920)$	$\Delta(1905)$
								$\Delta(1950)$	
2	S	$2^+$	$1/2$		$3/2^+$	$5/2^+$			
					$N(1720)$	$N(1680)$			
2	S	$0^+$	$3/2$				$3/2^+$		
							$\Delta(1600)$		
2	S	$0^+$	$1/2$	$1/2^+$					
				$N(1440)$					
1	M	$1^-$	$3/2$	$1/2^-$	$3/2^-$	$5/2^-$			
				$N(1650)$	$N(1700)$	$N(1675)$			
1	M	$1^-$	$1/2$	$1/2^-$	$3/2^-$		$1/2^-$	$3/2^-$	
				$N(1535)$	$N(1520)$		$\Delta(1620)$	$\Delta(1700)$	
0	S	$0^+$	$3/2$				$3/2^+$		
							$\Delta(1232)$		
0	S	$0^+$	$1/2$	$1/2^+$					
				$N(938)$					

up to the  $N=1$  band. Some more tentative assignments for higher multiplets are suggested in Ref. 30.

In Table 15.6, quark-model assignments are given for many of the established baryons whose  $SU(6) \otimes O(3)$  compositions are relatively unmixed. One must, however, keep in mind that apart from the mixing of the  $\Lambda$  singlet and octet states, states with same  $J^P$  but different  $L, S$  combinations can also mix. In the quark model with one-gluon exchange motivated interactions, the size of the mixing is determined by the relative strength of the tensor term with respect to the contact term (see below). The mixing is more important for the decay patterns of the states than for their positions. An example are the lowest lying  $(70, 1_1^-)$  states with  $J^P=1/2^-$  and  $3/2^-$ . The physical states are:

$$|N(1535)1/2^-\rangle = \cos(\Theta_S)|N^2P_M1/2^-\rangle - \sin(\Theta_S)|N^4P_M1/2^-\rangle \quad (15.26)$$

$$|N(1520)3/2^-\rangle = \cos(\Theta_D)|N^2P_M3/2^-\rangle - \sin(\Theta_D)|N^4P_M3/2^-\rangle \quad (15.27)$$

and the orthogonal combinations for  $N(1650)1/2^-$  and  $N(1700)3/2^-$ . The mixing is large for the  $J^P=1/2^-$  states ( $\Theta_S \approx -32^\circ$ ), but small for the  $J^P=3/2^-$  states ( $\Theta_D \approx +6^\circ$ ) [31,32].

All baryons of the ground state multiplets are known. Many of their properties, in particular their masses, are in good agreement even with the most basic versions of the quark model, including harmonic (or linear) confinement and a spin-spin interaction, which is responsible for the octet - decuplet mass shifts. A consistent description of the ground-state electroweak properties, however, requires refined relativistic constituent quark models.

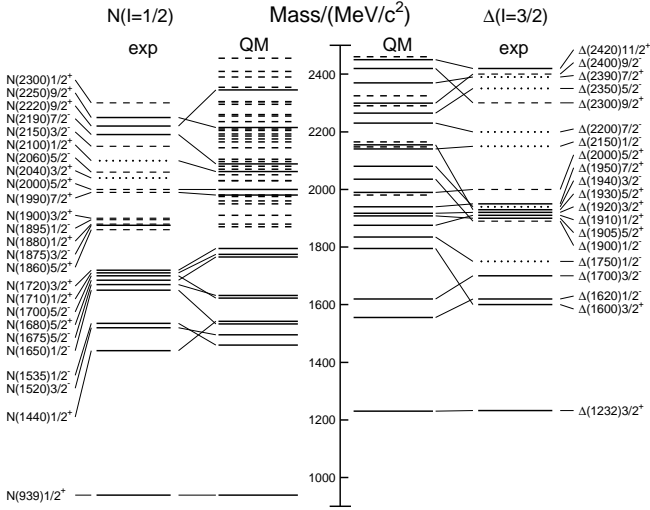
The situation for the excited states is much less clear. The assignment of some experimentally observed states with strange quarks to model configurations is only tentative and in many cases candidates are completely missing. Recently, Melde, Plessas and Sengl [33] have calculated baryon properties in relativistic constituent quark models, using one-gluon exchange and Goldstone-boson

exchange for the modeling of the hyperfine interactions (see Sec. 15.5 on Dynamics). Both types of models give qualitatively comparable results, and underestimate in general experimentally observed decay widths. Nevertheless, in particular on the basis of the observed decay patterns, the authors have assigned some additional states with strangeness to the  $SU(3)$  multiplets and suggest re-assignments for a few others. Among the new assignments are states with weak experimental evidence (two or three star ratings) and partly without firm spin/parity assignments, so that further experimental efforts are necessary before final conclusions can be drawn. We have added their suggestions in Table 15.6.

In the non-strange sector there are two main problems which are illustrated in Fig. 15.5, where the experimentally observed excitation spectrum of the nucleon ( $N$  and  $\Delta$  resonances) is compared to the results of a typical quark model calculation [34]. The lowest states from the  $N=2$  band, the  $N(1440)1/2^+$ , and the  $\Delta(1600)3/2^+$ , appear lower than the negative parity states from the  $N=1$  band (see Table 15.5) and much lower than predicted by most models. Also negative parity  $\Delta$  states from the  $N=3$  band ( $\Delta(1900)1/2^-$ ,  $\Delta(1940)3/2^-$ , and  $\Delta(1930)5/2^-$ ) are too low in energy. Part of the problem could be experimental. Among the negative parity  $\Delta$  states, only the  $\Delta(1930)5/2^-$  has three stars and the uncertainty in the position of the  $\Delta(1600)3/2^+$  is large (1550 - 1700 MeV).

Furthermore, many more states are predicted than observed. This has been known for a long time as the 'missing resonance' problem [31]. Up to an excitation energy of 2.4 GeV, about 45  $N$  states are predicted, but only 14 are established (four- or three-star; see Note on  $N$  and  $\Delta$  Resonances for the rating of the status of resonances) and 10 are tentative (two- or one-star). Even for the  $N=2$  band, up to now only half of the predicted states have been observed. The most recent partial wave analysis of elastic pion scattering and charge exchange data by Arndt and collaborators [35] has made the situation even worse. They found no evidence for almost half of the states listed in this review (and included in Fig. 15.5).

Such analyses are of course biased against resonances which couple only weakly to the  $N\pi$  channel. Quark model predictions for the couplings to other hadronic channels and to photons are given in Ref. 34. A large experimental effort is ongoing at several electron accelerators to study the baryon resonance spectrum with real and virtual photon-induced meson production reactions. This includes the search for as-yet-unobserved states, as well as detailed studies of the properties of the low lying states (decay patterns, electromagnetic couplings, magnetic moments, *etc.*) (see Ref. 36 for recent reviews). This experimental effort has currently entered its final phase with the measurement of single and double polarization observables for many different meson production channels, so that a much better understanding of the experimental spectrum can be expected for the near future.



**Figure 15.5:** Excitation spectrum of the nucleon. Compared are the positions of the excited states identified in experiment, to those predicted by a relativized quark model calculation. Left hand side: isospin  $I = 1/2$   $N$ -states, right hand side: isospin  $I = 3/2$   $\Delta$ -states. Experimental: (columns labeled 'exp'), three- and four-star states are indicated by full lines (two-star dashed lines, one-star dotted lines). At the very left and right of the figure, the spectroscopic notation of these states is given. Quark model [34]: (columns labeled 'QM'), all states for the  $N=1,2$  bands, low-lying states for the  $N=3,4,5$  bands. Full lines: at least tentative assignment to observed states, dashed lines: so far no observed counterparts. Many of the assignments between predicted and observed states are highly tentative.

In quark models, the number of excited states is determined by the effective degrees of freedom, while their ordering and decay properties are related to the residual quark - quark interaction. An overview of quark models for baryons is given in Ref. 32, recent discussions of baryon spectroscopy are given in Refs. 30 and 25. The effective degrees of freedom in the standard nonrelativistic quark model are three equivalent valence quarks with one-gluon exchange-motivated, flavor-independent color-magnetic interactions. The QCD aspect of gluon-gluon interactions is emphasized by the hypercentral quark model [37], [38], which includes in a natural way three-body forces between the quarks. A different class of models uses interactions which give rise to a quark - diquark clustering of the baryons: for a review see Ref. 39. If there is a tightly bound diquark, only two degrees of freedom are available at low energies, and thus *fewer* states are predicted. Furthermore, selection rules in the decay pattern may arise from the quantum numbers of the diquark. *More* states are predicted by collective models of the baryon like the algebraic approach in Ref. 40. In this approach, the quantum numbers of the valence quarks are distributed over a Y-shaped string-like configuration, and additional states arise *e.g.*, from vibrations of the strings. *More* states are also predicted in the framework of flux-tube models, see Ref. 41, which are motivated by lattice QCD. In addition to the quark

degrees of freedom, flux-tubes responsible for the confinement of the quarks are considered as degrees of freedom. These models include hybrid baryons containing explicit excitations of the gluon fields. However, since all half integral  $J^P$  quantum numbers are possible for ordinary baryons, such 'exotics' will be very hard to identify, and probably always mix with ordinary states. So far, the experimentally observed number of states is still far lower even than predicted by the quark-diquark models.

**Table 15.6:** Quark-model assignments for some of the known baryons in terms of a flavor-spin SU(6) basis. Only the dominant representation is listed. Assignments for several states, especially for the  $\Lambda(1810)$ ,  $\Lambda(2350)$ ,  $\Xi(1820)$ , and  $\Xi(2030)$ , are merely educated guesses. <sup>†</sup> recent suggestions for assignments and re-assignments from Ref. 33. For assignments of the charmed baryons, see the "Note on Charmed Baryons" in the Particle Listings.

$J^P$	$(D, L_N^P) S$	Octet members			Singlets
$1/2^+$	$(56, 0_0^+)$	$1/2 N(939)$	$\Lambda(1116)$	$\Sigma(1193)$	$\Xi(1318)$
$1/2^+$	$(56, 0_2^+)$	$1/2 N(1440)$	$\Lambda(1600)$	$\Sigma(1660)$	$\Xi(1690)^{\dagger}$
$1/2^-$	$(70, 1_1^-)$	$1/2 N(1535)$	$\Lambda(1670)$	$\Sigma(1620)$	$\Xi(?)$ $\Lambda(1405)$
				$\Sigma(1560)^{\dagger}$	
$3/2^-$	$(70, 1_1^-)$	$1/2 N(1520)$	$\Lambda(1690)$	$\Sigma(1670)$	$\Xi(1820)$ $\Lambda(1520)$
$1/2^-$	$(70, 1_1^-)$	$3/2 N(1650)$	$\Lambda(1800)$	$\Sigma(1750)$	$\Xi(?)$
				$\Sigma(1620)^{\dagger}$	
$3/2^-$	$(70, 1_1^-)$	$3/2 N(1700)$	$\Lambda(?)$	$\Sigma(1940)^{\dagger}$	$\Xi(?)$
$5/2^-$	$(70, 1_1^-)$	$3/2 N(1675)$	$\Lambda(1830)$	$\Sigma(1775)$	$\Xi(1950)^{\dagger}$
$1/2^+$	$(70, 0_2^+)$	$1/2 N(1710)$	$\Lambda(1810)$	$\Sigma(1880)$	$\Xi(?)$ $\Lambda(1810)^{\dagger}$
$3/2^+$	$(56, 2_2^+)$	$1/2 N(1720)$	$\Lambda(1890)$	$\Sigma(?)$	$\Xi(?)$
$5/2^+$	$(56, 2_2^+)$	$1/2 N(1680)$	$\Lambda(1820)$	$\Sigma(1915)$	$\Xi(2030)$
$7/2^-$	$(70, 3_3^-)$	$1/2 N(2190)$	$\Lambda(?)$	$\Sigma(?)$	$\Xi(?)$ $\Lambda(2100)$
$9/2^-$	$(70, 3_3^-)$	$3/2 N(2250)$	$\Lambda(?)$	$\Sigma(?)$	$\Xi(?)$
$9/2^+$	$(56, 4_4^+)$	$1/2 N(2220)$	$\Lambda(2350)$	$\Sigma(?)$	$\Xi(?)$
<hr/>					
Decuplet members					
$3/2^+$	$(56, 0_0^+)$	$3/2 \Delta(1232)$	$\Sigma(1385)$	$\Xi(1530)$	$\Omega(1672)$
$3/2^+$	$(56, 0_2^+)$	$3/2 \Delta(1600)$	$\Sigma(1690)^{\dagger}$	$\Xi(?)$	$\Omega(?)$
$1/2^-$	$(70, 1_1^-)$	$1/2 \Delta(1620)$	$\Sigma(1750)^{\dagger}$	$\Xi(?)$	$\Omega(?)$
$3/2^-$	$(70, 1_1^-)$	$1/2 \Delta(1700)$	$\Sigma(?)$	$\Xi(?)$	$\Omega(?)$
$5/2^+$	$(56, 2_2^+)$	$3/2 \Delta(1905)$	$\Sigma(?)$	$\Xi(?)$	$\Omega(?)$
$7/2^+$	$(56, 2_2^+)$	$3/2 \Delta(1950)$	$\Sigma(2030)$	$\Xi(?)$	$\Omega(?)$
$11/2^+$	$(56, 4_4^+)$	$3/2 \Delta(2420)$	$\Sigma(?)$	$\Xi(?)$	$\Omega(?)$

Recently, the influence of chiral symmetry on the excitation spectrum of the nucleon has been hotly debated from a somewhat new perspective. Chiral symmetry, the fundamental symmetry of QCD, is strongly broken for the low lying states, resulting in large mass differences of parity partners like the  $J^P=1/2^+$   $N(938)1/2^+$  ground state and the  $J^P=1/2^-$   $N(1535)1/2^-$  excitation. However, at higher excitation energies there is some evidence for parity doublets and even some very tentative suggestions for full chiral multiplets of  $N^*$  and  $\Delta$  resonances. An effective restoration of chiral symmetry at high excitation energies due to a decoupling from the quark condensate of the vacuum has been discussed (see Ref. 42 for recent reviews) as a possible cause. In this case, the mass generating mechanisms for low and high lying states would be essentially different. As a further consequence, the parity doublets would decouple from pions, so that experimental bias would be worse. However, parity doublets might also arise from the spin-orbital dynamics of the 3-quark system. Presently, the status of data does not allow final conclusions.

The most recent developments on the theory side are the first unquenched lattice calculations for the excitation spectrum discussed in Sec. 15.6. The results are basically consistent with the level counting of  $SU(6) \otimes O(3)$  in the standard non-relativistic quark model and show no indication for quark-diquark structures or parity

doubling. Consequently, there is as yet no indication from lattice that the mis-match between the excitation spectrum predicted by the standard quark model and experimental observations is due to inappropriate degrees of freedom in the quark model.

### 15.5. Dynamics

Quantum chromodynamics (QCD) is well-established as the theory for the strong interactions. As such, one of the goals of QCD is to predict the spectrum of strongly-interacting particles. To date, the only first-principles calculations of spectroscopy from QCD use lattice methods. These are the subject of Sec. 15.6. These calculations are difficult and unwieldy, and many interesting questions do not have a good lattice-based method of solution. Therefore, it is natural to build models, whose ingredients are abstracted from QCD, or from the low-energy limit of QCD (such as chiral Lagrangians) or from the data itself. The words “quark model” are a shorthand for such phenomenological models. Many specific quark models exist, but most contain a similar basic set of dynamical ingredients. These include:

- i) A confining interaction, which is generally spin-independent (*e.g.*, harmonic oscillator or linear confinement);
- ii) Different types of spin-dependent interactions:
  - a) commonly used is a color-magnetic flavor-independent interaction modeled after the effects of gluon exchange in QCD (see *e.g.*, Ref. 43) For example, in the  $S$ -wave states, there is a spin-spin hyperfine interaction of the form

$$H_{HF} = -\alpha_S M \sum_{i>j} (\vec{\sigma} \lambda_a)_i (\vec{\sigma} \lambda_a)_j, \quad (15.28)$$

where  $M$  is a constant with units of energy,  $\lambda_a$  ( $a = 1, \dots, 8$ ) is the set of SU(3) unitary spin matrices, defined in Sec. 45, on “SU(3) Isoscalar Factors and Representation Matrices,” and the sum runs over constituent quarks or antiquarks. Spin-orbit interactions, although allowed, seem to be small in general, but a tensor term is responsible for the mixing of states with the same  $J^P$  but different  $L, S$  combinations.

b) other approaches include flavor-dependent short-range quark forces from instanton effects (see *e.g.*, Ref. 44) This interaction acts only on scalar, isoscalar pairs of quarks in a relative  $S$ -wave state:

$$\langle q^2; S, L, T | W | q^2; S, L, T \rangle = -4g\delta_{S,0}\delta_{L,0}\delta_{I,0}\mathcal{W} \quad (15.29)$$

where  $\mathcal{W}$  is the radial matrix element of the contact interaction.

c) a rather different and controversially discussed approach is based on flavor-dependent spin-spin forces arising from one-boson exchange. The interaction term is of the form:

$$H_{HF} \propto \sum_{i<j} V(\vec{r}_{ij}) \lambda_i^F \cdot \lambda_j^F \vec{\sigma}_i \cdot \vec{\sigma}_j \quad (15.30)$$

where the  $\lambda_i^F$  are in flavor space (see *e.g.*, Ref. 45).

- iii) A strange quark mass somewhat larger than the up and down quark masses, in order to split the SU(3) multiplets;
- iv) In the case of spin-spin interactions (iia,c), a flavor-symmetric interaction for mixing  $q\bar{q}$  configurations of different flavors (*e.g.*,  $u\bar{u} \leftrightarrow d\bar{d} \leftrightarrow s\bar{s}$ ), in isoscalar channels, so as to reproduce *e.g.*, the  $\eta - \eta'$  and  $\omega - \phi$  mesons.

These ingredients provide the basic mechanisms that determine the hadron spectrum in the standard quark model.

### 15.6. Lattice Calculations of Hadronic Spectroscopy

Lattice calculations are a major source of information about QCD masses and matrix elements. The necessary theoretical background is given in Sec. 18 of this *Review*. Here we confine ourselves to some general comments and illustrations of lattice calculations for spectroscopy.

In general, the cleanest lattice results come from computations of processes in which there is only one particle in the simulation

volume. These quantities include masses of hadrons, simple decay constants, like pseudoscalar meson decay constants, and semileptonic form factors (such as the ones appropriate to  $B \rightarrow D\ell\nu, K\ell\nu, \pi\ell\nu$ ). The cleanest predictions for masses are for states which have narrow decay widths and are far below any thresholds to open channels, since the effects of final state interactions are not yet under complete control on the lattice. As a simple corollary, the lightest state in a channel is easier to study than the heavier ones. “Difficult” states for the quark model (such as exotics) are also difficult for the lattice because of the lack of simple operators which couple well to them.

Good-quality modern lattice calculations will present multi-part error budgets with their predictions. A small part of the uncertainty is statistical, from sample size. Typically, the quoted statistical uncertainty includes uncertainty from a fit: it is rare that a simulation computes one global quantity which is the desired observable. Simulations which include virtual quark-antiquark pairs (also known as “dynamical quarks” or “sea quarks”) are often done at up and down quark mass values heavier than the experimental ones, and it is then necessary to extrapolate in these quark masses. Simulations can work at the physical values of the heavier quarks’ masses. They are always done at nonzero lattice spacing, and so it is necessary to extrapolate to zero lattice spacing. Some theoretical input is needed to do this. Much of the uncertainty in these extrapolations is systematic, from the choice of fitting function. Other systematics include the effect of finite simulation volume, the number of flavors of dynamical quarks actually simulated, and technical issues with how these dynamical quarks are included. The particular choice of a fiducial mass (to normalize other predictions) is not standardized; there are many possible choices, each with its own set of strengths and weaknesses, and determining it usually requires a second lattice simulation from that used to calculate the quantity under consideration.

A systematic error of major historical interest is the “quenched approximation,” in which dynamical quarks are simply left out of the simulation. This was done because the addition of these virtual pairs presented an expensive computational problem. No generally-accepted methodology has ever allowed one to correct for quenching effects, short of redoing all calculations with dynamical quarks. Recent advances in algorithms and computer hardware have rendered it obsolete.

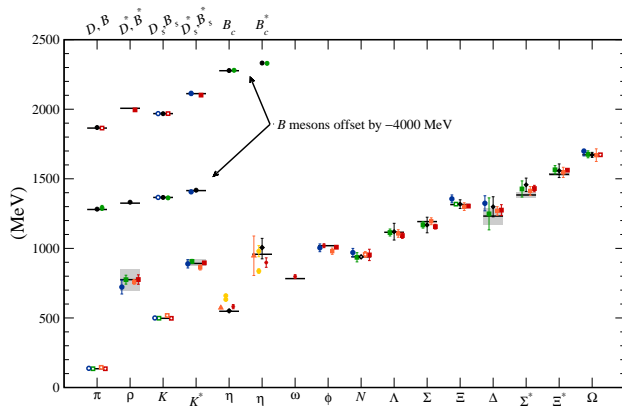
With these brief remarks, we turn to examples. The field of lattice QCD simulations is vast, and so it is not possible to give a comprehensive review of them in a small space. The history of lattice QCD simulations is a story of thirty years of incremental improvements in physical understanding, algorithm development, and ever faster computers, which have combined to bring the field to a present state where it is possible to carry out very high quality calculations. We present a few representative illustrations, to show the current state of the art.

By far, the major part of all lattice spectroscopy is concerned with that of the light hadrons, and so we illustrate results in Fig. 15.6, a comprehensive summary provided by A. Kronfeld [46].

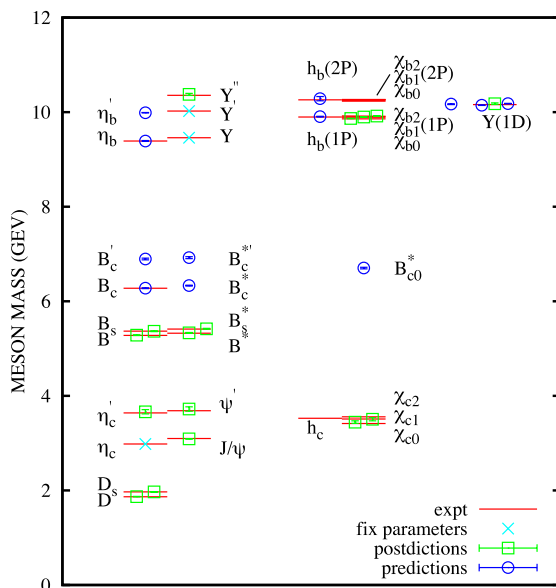
Flavor singlet mesons are at the frontier of lattice QCD calculations, because one must include the effects of “annihilation graphs,” for the valence  $q$  and  $\bar{q}$ . Recently, several groups, Refs. 6, 53–56, have reported calculations of the  $\eta$  and  $\eta'$  mesons. The numbers of Ref. 6 are typical, finding masses of 573(6) and 947(142) MeV for the  $\eta$  and  $\eta'$ . The singlet-octet mixing angle (in the conventions of Table 15.2) is  $\theta_{in} = -14.1(2.8)^\circ$ .

The spectroscopy of mesons containing heavy quarks has become a truly high-precision endeavor. These simulations use Non-Relativistic QCD (NRQCD) or Heavy Quark Effective Theory (HQET), systematic expansions of the QCD Lagrangian in powers of the heavy quark velocity, or the heavy quark mass. Terms in the Lagrangian have obvious quark model analogs, but are derived directly from QCD. For example, the heavy quark potential is a derived quantity, extracted from simulations. Fig. 15.7 shows the mass spectrum for mesons containing at least one heavy ( $b$  or  $c$ ) quark from Ref. 59 It also contains results from Refs. 61 and 62 The calculations uses a discretization of nonrelativistic QCD for bottom quarks with charm and lighter quarks being handled with an improved relativistic action. Four flavors (u, d, s, c) of dynamical quarks are included.





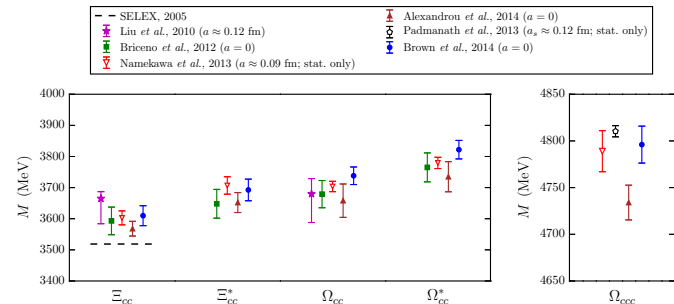
**Figure 15.6:** Hadron spectrum from lattice QCD. Comprehensive results for mesons and baryons are from MILC [47,48], PACS-CS [49], BMW [50], QCDSF [51], and ETM [68]. Results for  $\eta$  and  $\eta'$  are from RBC & UKQCD [6], Hadron Spectrum [54] (also the only  $\omega$  mass), UKQCD [53], and Michael, Otnad, and Urbach [55]. Results for heavy-light hadrons from Fermilab-MILC [57], HPQCD [58,59], and Mohler and Woloshyn [60]. Circles, squares, diamonds, and triangles stand for staggered, Wilson, twisted-mass Wilson, and chiral sea quarks, respectively. Asterisks represent anisotropic lattices. Open symbols denote the masses used to fix parameters. Filled symbols (and asterisks) denote results. Red, orange, yellow, green, and blue stand for increasing numbers of ensembles (i.e., lattice spacing and sea quark mass). Black symbols stand for results with 2+1+1 flavors of sea quarks. Horizontal bars (gray boxes) denote experimentally measured masses (widths).  $b$ -flavored meson masses are offset by  $-4000$  MeV.



**Figure 15.7:** Spectroscopy for mesonic systems containing one or more heavy quarks (adapted from Ref. 59) Particles whose masses are used to fix lattice parameters are shown with crosses; the authors distinguish between “predictions” and “postdictions” of their calculation. Lines represent experiment.

Fig. 15.8 shows a compilation of recent lattice results for doubly and triply charmed baryons, from Ref. 63.

There are a number of reported states near the charmonium- $D - \bar{D}$  threshold, including the  $X(3872)$ ,  $D_{s0}^*(2317)$ ,  $Z_c^\pm(3900)$  and  $X(4140)$ , whose quark composition is obscure (see the “Note on Non- $q\bar{q}$  mesons in the Meson Listings.” The current status of lattice studies of these



**Figure 15.8:** Comparison of lattice QCD results for the doubly and triply charmed baryon masses. Labels are Liu, *et al.*, [64]; Briceño, *et al.*, [65]; Namekawa, *et al.*, [66]; Padmanath, *et al.*, [67]; Alexandrou, *et al.*, [68]; and Brown, *et al.*, citequarkmod:Brown:2014ena. Only calculations with dynamical light quarks are included; for the doubly charmed baryons, only calculations were performed at or extrapolated to the physical pion mass are shown. Results without estimates of systematic uncertainties are labeled “stat. only”. The lattice spacing values used in the calculations are also given;  $a = 0$  indicates that the results have been extrapolated to the continuum limit. In the plot of the doubly charmed baryons, the unconfirmed experimental result for the  $\Xi_{cc}^+$  mass from SELEX [26,27] is shown with a dashed line.

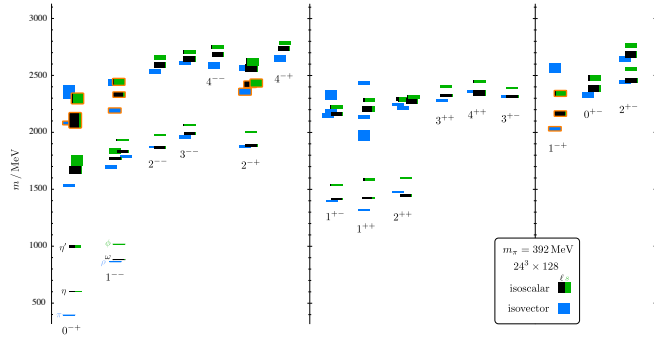
states is reviewed in Ref. 69.

Recall that lattice calculations take operators which are interpolating fields with quantum numbers appropriate to the desired states, compute correlation functions of these operators, and fit the correlation functions to functional forms parametrized by a set of masses and matrix elements. As we move away from hadrons which can be created by the simplest quark model operators (appropriate to the lightest meson and baryon multiplets) we encounter a host of new problems: either no good interpolating fields, or too many possible interpolating fields, and many states with the same quantum numbers. Techniques for dealing with these interrelated problems vary from collaboration to collaboration, but all share common features: typically, correlation functions from many different interpolating fields are used, and the signal is extracted in what amounts to a variational calculation using the chosen operator basis. In addition to mass spectra, wave function information can be garnered from the form of the best variational wave function. Of course, the same problems which are present in the spectroscopy of the lightest hadrons (the need to extrapolate to infinite volume, physical values of the light quark masses, and zero lattice spacing) are also present. We briefly touch on three different kinds of hadrons: excited states of mesons (including hybrids), excited states of baryons, and glueballs. The quality of the data is not as good as for the ground states, and so the results continue to evolve.

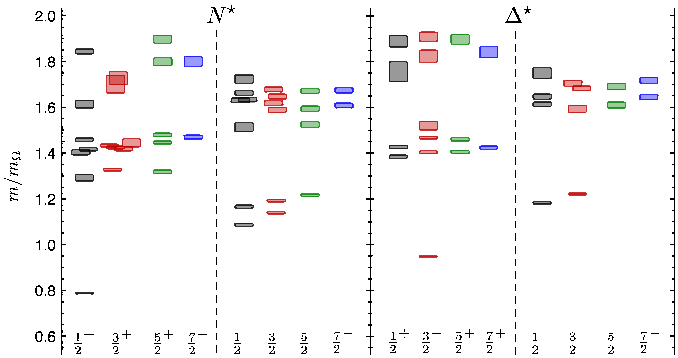
Modern calculations use a large bases of trial states, which allow them to probe many quantum number channels simultaneously. This is vital for studying “difficult sectors” of QCD, such as the isoscalar mesons. A recent example of meson spectroscopy where this is done, by Ref. 56, is shown in Fig. 15.9. The quark masses are still heavier than their physical values, so the pion is at 391 MeV. The authors can assign a relative composition of nonstrange and strange quark content to their states, observing, for example, a nonstrange  $\omega$  and a strange  $\phi$ . Some states also have a substantial component of gluonic excitation. Note especially the three exotic channels  $J^{PC} = 1^{-+}$ ,  $0^{+-}$ , and  $2^{+-}$ , with states around 2 GeV. These calculations will continue to improve as the quark masses are carried lower.

The interesting physics questions of excited baryon spectroscopy to be addressed are precisely those enumerated in the last section. An example of a recent calculation, due to Ref. 70 is shown in Fig. 15.10. Notice that the pion is not yet at its physical value. The lightest positive parity state is the nucleon, and the Roper resonance has not yet appeared as a light state.

In Fig. 15.3 we showed a figure from Ref. 11 presenting a lattice prediction for the glueball mass spectrum in quenched approximation. A true QCD prediction of the glueball spectrum requires dynamical



**Figure 15.9:** Isoscalar (green and black) and isovector (blue) spectrum from Ref. 56. States are labeled  $J^{PC}$ . The quark mass is heavier than its physical value;  $m_\pi = 391$  MeV. The vertical height of each box indicates the statistical uncertainty in the mass. Black and green indicate relative nonstrange and strange composition. Orange outlines show states with a large chromomagnetic component to their wave function, which the authors of Ref. 56 argue are hybrid states. Note the exotic states in the three rightmost columns.

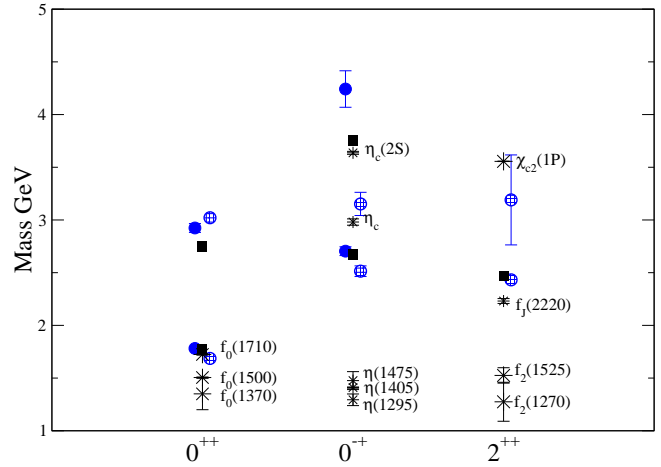


**Figure 15.10:** Spin-identified spectrum of nucleons and deltas, from lattices where  $m_\pi = 396$  MeV, in units of the calculated  $\Omega$  mass, from Ref. 70. The colors just correspond to the different  $J$  assignments: grey for  $J = 1/2$ , red for  $J = 3/2$ , green for  $J = 5/2$ , blue for  $J = 7/2$ .

light quarks and (because glueball operators are intrinsically noisy) high statistics. Only recently have the first useful such calculations appeared. Fig. 15.11 shows results from Ref. 71, done with dynamical  $u$ ,  $d$  and  $s$  quarks at two lattice spacings, 0.123 and 0.092 fm, along with comparisons to the quenched lattice calculation of Ref. 10 and to experimental isosinglet mesons. The dynamical simulation is, of course, not the last word on this subject, but it shows that the effects of quenching seem to be small.

Several other features of hadronic spectroscopy are also being studied on the lattice.

Electromagnetic mass splittings (such as the neutron - proton mass difference) are interesting but difficult. These calculations are important for determining the values of the quark masses (for a discussion see the review in the PDG). Knowing that the neutron is heavier than the proton tells us that these splittings have a complicated origin. One part of the shift is because the up and down quarks have slightly different masses. The second is that the quarks have (different) charges. In pre-lattice days, phenomenologists would combine Coulomb forces and spin-dependent electromagnetic hyperfine interactions to model their charge effects. These days, in order to compute hadronic mass differences on the lattice, electromagnetic interactions must be included in the simulations. This creates a host of technical issues. An important one is that electromagnetic interactions are long range, but lattice simulations are done in finite volumes. A recent calculation, Ref. 72, has presented the first results for electromagnetic mass splittings in the baryon octet. The situation is summarized in the review Ref. 73.



**Figure 15.11:** Lattice QCD predictions for glueball masses. The open and closed circles are the larger and smaller lattice spacing data of the full QCD calculation of glueball masses of Ref. 71. Squares are the quenched data for glueball masses of Ref. 10. The bursts labeled by particle names are experimental states with the appropriate quantum numbers.

Most hadrons are resonances, and their widths are the last target of lattice simulations we will mention. The actual calculation is of the combined mass of two (or more) hadrons in a box of finite size. The combined mass is shifted from being the sum of the individual masses because the finite box forces the hadrons to interact with each other. The volume-dependent mass shift yields the phase shift for the continuum scattering amplitude, which in turn can be used to extract the resonance mass and width, with some degree of modeling. So far only two-body resonances, the rho meson and a few others, have been well studied. This is an active research topic. A recent review, Ref. 74, summarizes the situation, and an example of a calculation of the rho meson's decay width is Ref. 75.

## References:

1. J. Schwinger, Phys. Rev. **135**, B116 (1964).
2. A. Bramon *et al.*, Phys. Lett. **B403**, 339 (1997).
3. A. Aloisio *et al.*, Phys. Lett. **B541**, 45 (2002).
4. C. Amsler *et al.*, Phys. Lett. **B294**, 451 (1992).
5. C. Amsler, Rev. Mod. Phys. **70**, 1293 (1998).
6. N.H. Christ *et al.*, Phys. Rev. Lett. **105**, 241601 (2010).
7. T. Feldmann, Int. J. Mod. Phys. **A915**, 159 (2000).
8. C. Amsler and F.E. Close, Phys. Rev. **D53**, 295 (1996).
9. R.L. Jaffe, Phys. Rev. **D15**, 267 (1977); R.L. Jaffe, Phys. Rev. **D15**, 281 (1977).
10. C. Morningstar and M. Peardon, Phys. Rev. **D60**, 034509 (1999).
11. Y. Chen *et al.*, Phys. Rev. **D73**, 014516 (2006).
12. W.J. Lee and D. Weingarten, Phys. Rev. **D61**, 014015 (2000).
13. G.S. Bali *et al.*, Phys. Lett. **B309**, 378 (1993).
14. C. Michael, AIP Conf. Proc. **432**, 657 (1998).
15. F.E. Close and A. Kirk, Eur. Phys. J. **C21**, 531 (2001).
16. W. Ochs, J. Phys. **G40**, 043001 (2013).
17. C. Amsler and N.A. Törnqvist, Phys. Reports **389**, 61 (2004).
18. N. Isgur and J. Paton, Phys. Rev. **D31**, 2910 (1985).
19. P. Lacock *et al.*, Phys. Lett. **B401**, 308 (1997); C. Bernard *et al.*, Phys. Rev. **D56**, 7039 (1997); C. Bernard *et al.*, Phys. Rev. **D68**, 074505 (2003).
20. M. Chanowitz and S. Sharpe, Nucl. Phys. **B222**, 211 (1983).
21. T. Barnes *et al.*, Nucl. Phys. **B224**, 241 (1983).
22. W.-M. Yao *et al.*, J. Phys. **G33**, 1 (2006).
23. R. Aaij *et al.*, Phys. Rev. Lett. **115**, 072001 (2015).
24. F.E. Close, in *Quarks and Nuclear Forces* (Springer-Verlag, 1982), p. 56.
25. V. Crede and W. Roberts, Rept. on Prog. in Phys. **76**, 076301 (2013).
26. M. Mattson *et al.*, Phys. Rev. Lett. **89**, 112001 (2002).



27. A. Ocherashvili *et al.*, Phys. Lett. **B628**, 18 (2005).
28. M. Karliner and J.L. Rosner, Phys. Rev. **D90**, 094007 (2014).
29. R.H. Dalitz and L.J. Reinders, in “Hadron Structure as Known from Electromagnetic and Strong Interactions,” *Proceedings of the Hadron ’77 Conference* (Veda, 1979), p. 11.
30. E. Klempt and J.M. Richard, Rev. Mod. Phys. **82**, 1095 (2010).
31. N. Isgur and G. Karl, Phys. Rev. **D18**, 4187 (1978); *ibid.*, **D19**, 2653 (1979); *ibid.*, **D20**, 1191 (1979).
32. S. Capstick and W. Roberts, Prog. in Part. Nucl. Phys. **45**, 241 (2000).
33. T. Melde, W. Plessas, and B. Sengl, Phys. Rev. **D77**, 114002 (2008).
34. S. Capstick and W. Roberts, Phys. Rev. **D49**, 4570 (1994); *ibid.*, **D57**, 4301 (1998); *ibid.*, **D58**, 074011 (1998); S. Capstick, Phys. Rev. **D46**, 2864 (1992).
35. R.A. Arndt *et al.*, Phys. Rev. **C74**, 045205 (2006).
36. B. Krusche and S. Schadmand, Prog. in Part. Nucl. Phys. **51**, 399 (2003); V.D. Burkert and T.-S.H. Lee, Int. J. Mod. Phys. **E13**, 1035 (2004); see also A.J.G. Hey and R.L. Kelly, Phys. Reports **96**, 71 (1983).
37. M. Ferraris *et al.*, Phys. Lett. **B364**, 231 (1995).
38. M.M. Giannini and E. Santopinto, Chin. J. Phys. **53**, 020301-1 (2015).
39. M. Anselmino *et al.*, Rev. Mod. Phys. **65**, 1199 (1993).
40. R. Bijker *et al.*, Ann. Phys. **236**, 69 (1994).
41. N. Isgur and J. Paton, Phys. Rev. **D31**, 2910 (1985); S. Capstick and P.R. Page, Phys. Rev. **C66**, 065204 (2002).
42. R.L. Jaffe, D. Pirjol, and A. Scardicchio, Phys. Rev. **435**, 157 (2006); L. Ya. Glozman, Phys. Rev. **444**, 1 (2007).
43. A. De Rujula *et al.*, Phys. Rev. **D12**, 147 (1975).
44. W.H. Blask *et al.*, Z. Phys. **A337**, 327 (1990); U. Löring *et al.*, Eur. Phys. J. **A10**, 309 (2001); U. Löring *et al.*, Eur. Phys. J. **A10**, 395 (2001); *ibid.*, **A10**, 447 (2001).
45. L.Y. Glozman and D.O. Riska, Phys. Rev. **268**, 263 (1996); L.Y. Glozman *et al.*, Phys. Rev. **D58**, 094030 (1998); W. Plessas, Int. J. Mod. Phys. **A30**, 1530013 (2015).
46. A. Kronfeld, private communication. See also Ann. Rev. Nucl. and Part. Sci. **62**, 265 (2012) for an earlier version of this figure.
47. C. Aubin *et al.*, Phys. Rev. **D70**, 094505 (2004) [[hep-lat/0402030](#)].
48. A. Bazavov *et al.*, Rev. Mod. Phys. **82**, 1349 (2010) [[arXiv:0903.3598](#)].
49. S. Aoki *et al.* [PACS-CS Collab.], Phys. Rev. **D79**, 034503 (2009) [[arXiv:0807.1661](#)].
50. S. Durr *et al.*, Science **322**, 1224 (2008) [[arXiv:0906.3599](#)].
51. W. Bietenholz *et al.*, Phys. Rev. **D84**, 054509 (2011) [[arXiv:1102.5300](#)].
52. C. Alexandrou *et al.*, Phys. Rev. **D90**, 074501 (2014) [[arXiv:1406.4310](#)].
53. E.B. Gregory *et al.* [UKQCD Collab.], Phys. Rev. **D86**, 014504 (2012) [[arXiv:1112.4384](#)].
54. J.J. Dudek *et al.*, Phys. Rev. **D83**, 111502 (2011) [[arXiv:1102.4299](#)].
55. C. Michael *et al.* [ETM Collab.], Phys. Rev. Lett. **111**, 181602 (2013) [[arXiv:1310.1207](#)].
56. J.J. Dudek *et al.* [Hadron Spectrum Collab.], Phys. Rev. **D88**, 094505 (2013) [[arXiv:1309.2608](#)].
57. C. Bernard *et al.* [Fermilab Lattice and MILC Collab.], Phys. Rev. **D83**, 034503 (2011) [[arXiv:1003.1937](#)].
58. E.B. Gregory *et al.*, Phys. Rev. **D83**, 014506 (2011) [[arXiv:1010.3848](#)].
59. R.J. Dowdall *et al.*, Phys. Rev. **D86**, 094510 (2012) [[arXiv:1207.5149](#)].
60. D. Mohler and R. M. Woloshyn, Phys. Rev. **D84**, 054505 (2011) [[arXiv:1103.5506](#)].
61. J.O. Daldrop *et al.* [HPQCD Collab.], Phys. Rev. Lett. **108**, 102003 (2012).
62. G.C. Donald *et al.*, Phys. Rev. **D86**, 094501 (2012).
63. Z.S. Brown *et al.*, Phys. Rev. **D90**, 094507 (2014).
64. L. Liu *et al.*, Phys. Rev. **D81**, 094505 (2010).
65. R.A. Briceno, H.-W. Lin, and D.R. Bolton, Phys. Rev. **D86**, 094504 (2012).
66. Y. Namekawa *et al.* [PACS-CS Collab.], Phys. Rev. **D87**, 094512 (2013).
67. M. Padmanath *et al.*, Phys. Rev. **D90**, 074504 (2014).
68. C. Alexandrou *et al.*, Phys. Rev. **D90**, 074501 (2014).
69. S. Prelovsek, PoS LATTICE **2014**, 015 (2014) [[arXiv:1411.0405](#)].
70. R.G. Edwards *et al.*, Phys. Rev. **D84**, 074508 (2011) [[arXiv:1104.5152](#)].
71. C.M. Richards *et al.*, [UKQCD Collab.], Phys. Rev. **D82**, 034501 (2010).
72. S. Borsanyi *et al.*, Science **347**, 1452 (2015) [[arXiv:1406.4088](#)].
73. A. Portelli, PoS LATTICE **2014**, 013 (2015) [[arXiv:1505.07057](#)].
74. M. Döring, PoS LATTICE **2013**, 006 (2014)..
75. J.J. Dudek *et al.* [Hadron Spectrum Collab.], Phys. Rev. **D87**, 034505 (2013) [Phys. Rev. **D90**, 099902 (2014)] [[arXiv:1212.0830](#)].

## 16. GRAND UNIFIED THEORIES

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### 16.1. The Standard Model

The Standard Model (SM) may be defined as the renormalizable field theory with gauge group  $G_{SM} = SU(3)_C \times SU(2)_L \times U(1)_Y$ , with 3 generations of fermions in the representation

$$(\mathbf{3}, \mathbf{2})_{1/3} + (\bar{\mathbf{3}}, \mathbf{1})_{-4/3} + (\bar{\mathbf{3}}, \mathbf{1})_{2/3} + (\mathbf{1}, \mathbf{2})_{-1} + (\mathbf{1}, \mathbf{1})_2, \quad (16.1)$$

and a scalar Higgs doublet  $H$  transforming as  $(\mathbf{1}, \mathbf{2})_1$ . Here and below we use boldface numbers to specify the dimension of representations of non-Abelian groups (in this case fundamental and antifundamental) and lower indices for  $U(1)$  charges. The fields of Eq. (16.1) should also be familiar as  $[Q, u^c, d^c, L, e^c]$ , with  $Q = (u, d)$  and  $L = (\nu, e)$  being the quark and lepton  $SU(2)$ -doublets and  $u^c, d^c, e^c$  charge conjugate  $SU(2)$ -singlets.<sup>†</sup> Especially after the recent discovery of the Higgs, this model is remarkably complete and consistent with almost all experimental data.

A notable exception are neutrino masses, which are known to be non-zero but are absent in the SM even after the Higgs acquires its vacuum expectation value (VEV). The minimalist attitude is to allow for the dimension-five operator  $(HL)^2$ , which induces (Majorana) neutrino masses. In the seesaw mechanism [1,2,3] this operator is generated by integrating out heavy singlet fermions (r.h. neutrinos). Alternatively, neutrinos can have Dirac masses if light singlet neutrinos are added to the SM spectrum.

Conceptual problems of the SM include the absence of a Dark Matter candidate, of a mechanism for generating the baryon asymmetry of the universe, and of any reason for the observed smallness of the  $\theta$  parameter of QCD ( $\theta_{QCD}$ ). In addition, the apparently rather complex group-theoretic data of Eq. (16.1) remains unexplained. Together with the abundance of seemingly arbitrary coupling constants, this disfavors the SM as a candidate fundamental theory, even before quantum gravity problems arise at energies near  $M_P$ .

To be precise, there are 19 SM parameters which have to be fitted to data: Three gauge couplings\*  $g_3, g_2$  and  $g_1$ , 13 parameters associated with the Yukawa couplings (9 charged fermion masses, three mixing angles and one CP phase in the CKM matrix.), the Higgs mass and quartic coupling, and  $\theta_{QCD}$ . In addition, Majorana neutrinos introduce 3 more masses and 6 mixing angles and phases.

As we will see, the paradigm of grand unification addresses mainly the group theoretic data of Eq. (16.1) and the values of the three gauge couplings. In many concrete realizations, it then impacts also the other mentioned issues of the SM, such as e.g. the family structure and fermion mass hierarchy.

More specifically, after precision measurements of the Weinberg angle  $\theta_W$  in the LEP experiments, supersymmetric GUTs (SUSY GUTs) have become the leading candidates in the search for ‘Physics beyond the SM’. Supersymmetry (SUSY) is a symmetry between bosons and fermions which requires the addition of superpartners to the SM spectrum, thereby leading to the noted prediction of  $\theta_W$  [4]. The measured Higgs mass ( $\sim 125$  GeV) is in principle consistent with this picture, assuming superpartners in the region of roughly 10 TeV. Such heavy superpartners then induce radiative corrections raising the Higgs mass above the  $Z$  boson mass  $m_Z$  [5,6]. However, if SUSY is motivated as a solution to the gauge hierarchy problem (i.e. to the naturalness problem of the Higgs mass) [7], its minimal incarnation in terms of the MSSM is becoming questionable. Indeed, compared to expectations based on the **minimal** SUSY SM (MSSM) with superpartner masses below about 1 TeV, the Higgs mass is somewhat too high [8]. Independently, the LHC has disfavored light colored superpartners. These facts represent new hints for future work on SUSY GUTs or on GUTs without TeV-scale supersymmetry.

<sup>†</sup> In our convention the electric charge is  $Q = T_3 + Y/2$  and all our spinor fields are left-handed.

\* Equivalently, the  $SU(2)_L$  and  $U(1)_Y$  couplings are denoted as  $g = g_2$  and  $g' = \sqrt{3/5} g_1$ . One also uses  $\alpha_s = \alpha_3 = (g_3^2/4\pi)$ ,  $\alpha_{EM} = (e^2/4\pi)$  with  $e = g \sin \theta_W$  and  $\sin^2 \theta_W = (g')^2/(g^2 + (g')^2)$ .

### 16.2. Basic Group Theory and Charge Quantization

Historically, the first attempt at unification was the Pati-Salam model with gauge group  $G_{PS} = SU(4)_C \times SU(2)_L \times SU(2)_R$  [9]. It unifies SM fermions in the sense that one generation (plus an extra SM singlet) now comes from the  $(\mathbf{4}, \mathbf{2}, \mathbf{1}) + (\bar{\mathbf{4}}, \mathbf{1}, \mathbf{2})$  of  $G_{PS}$ . This is easy to verify from the breaking pattern  $SU(4)_C \rightarrow SU(3)_C \times U(1)_{B-L}$  together with the identification of SM hypercharge as a linear combination between  $B - L$  (baryon minus lepton number) with the  $T_3$  generator of  $SU(2)_R$ . This model explains charge quantization, that is, why all electric charges are integer multiples of some smallest charge in the SM. However,  $G_{PS}$  is not simple (containing three simple factors), and thus it does not predict gauge coupling unification.

Since  $G_{SM}$  has rank four (two for  $SU(3)_C$  and one for  $SU(2)_L$  and  $U(1)_Y$ , respectively), the rank-four group  $SU(5)$  is the minimal choice for unification in a simple group [10]. The three SM gauge coupling constants derive from a universal coupling  $\alpha_G$  at the GUT scale  $M_G$ . Explicitly embedding  $G_{SM}$  in  $SU(5)$  is straightforward, with  $SU(3)_C$  and  $SU(2)_L$  corresponding e.g. to the upper-left  $3 \times 3$  and lower-right  $2 \times 2$  blocks, respectively, in traceless  $5 \times 5$  matrices for  $SU(5)$  generators of the fundamental representation. The  $U(1)_Y$  corresponds to matrices generated by  $\text{diag}(-2/3, -2/3, -2/3, 1, 1)$ , which hence commute with  $SU(3)_C \times SU(2)_L \subset SU(5)$ . It is then easy to derive how one SM generation precisely comes from the  $\mathbf{10} + \bar{\mathbf{5}}$  of  $SU(5)$  (where  $\mathbf{10}$  is the antisymmetric rank-2 tensor):

$$\mathbf{10} : \begin{pmatrix} 0 & u_b^c & -u_g^c & u_r & d_r \\ -u_b^c & 0 & u_l^c & u_g & d_g \\ u_g^c & -u_r^c & 0 & u_b & d_b \\ -u_r & -u_g & -u_b & 0 & e^c \\ -d_r & -d_g & -d_b & -e^c & 0 \end{pmatrix} \quad \text{and} \quad \bar{\mathbf{5}} : \begin{pmatrix} d_r^c \\ d_g^c \\ d_b^c \\ e \\ -\nu_e \end{pmatrix} \quad (16.2)$$

Since  $SU(5)$  has 24 generators,  $SU(5)$  GUTs have 12 new gauge bosons known as  $X$  bosons (or  $X/Y$  bosons) in addition to the SM.  $X$  bosons form an  $SU(3)_C$ -triplet and  $SU(2)_L$ -doublet. Their interaction connects quarks and leptons such that baryon and lepton numbers are not conserved and nucleon decay is predicted. Furthermore,  $U(1)_Y$  hypercharge is automatically quantized since it is embedded in  $SU(5)$ .

In order to break the electroweak symmetry at the weak scale and give mass to quarks and leptons, Higgs doublets are needed. In the minimal  $SU(5)$  model, they can sit in either a  $\mathbf{5}_H$  or  $\bar{\mathbf{5}}_H$ . The three additional states are referred to as color-triplet Higgs scalars. Their couplings also violate baryon and lepton numbers, inducing nucleon decay. In order not to violently disagree with the non-observation of nucleon decay, the triplet mass must be greater than  $\sim 10^{11}$  GeV [11]. Moreover, in SUSY GUTs [12], in order to cancel anomalies as well as give mass to both up and down quarks, both Higgs multiplets  $\mathbf{5}_H$  and  $\bar{\mathbf{5}}_H$  are required. As we shall discuss later, nucleon decay now constrains the Higgs triplets to have mass significantly greater than  $M_G$  in the minimal SUSY  $SU(5)$  GUT since integrating out the Higgs triplets generates dimension-five baryon-number-violating operators [13]. The mass splitting between doublet and triplet in the  $\mathbf{5}_H$  (and  $\bar{\mathbf{5}}_H$ ) comes from their interaction with the  $SU(5)$  breaking sector.

While  $SU(5)$  allows for the minimal GUT models, unification is not complete: Two independent representations,  $\mathbf{10}$  and  $\bar{\mathbf{5}}$ , are required for one SM generation.

A further representation, an  $SU(5)$  singlet, has to be added to serve as r.h. neutrino in the seesaw mechanism. In this case, the r.h. neutrino masses are not necessarily related to the GUT scale.

By contrast, a single **16**-dimensional spinor representation of  $SO(10)$  accommodates a full SM generation together with an extra singlet, potentially providing a r.h. neutrino [14]. This is most easily understood from the breaking pattern  $SO(10) \rightarrow SU(5) \times U(1)_X$  and the associated branching rule\*  $\mathbf{16} = \mathbf{10}_{-5} + \bar{\mathbf{5}}_3 + \mathbf{1}_{-1}$ . Here the indices refer to charges under the  $U(1)_X$  subgroup, which is orthogonal to  $SU(5)$  and reflects the fact that  $SO(10)$  has rank five. From the above, it is easy to see that  $U(1)_X$  charges can be given as  $2Y - 5(B - L)$ .

\* Useful references on group theory in the present context include [15] and refs. therein.

Intriguingly, all representations of  $SO(10)$  are anomaly free in four dimensions (4d). Thus, the absence of anomalies in an  $SU(5)$ -GUT or a SM generation can be viewed as deriving from this feature.

Table 16.1 presents the states of one family of quarks and leptons, as they appear in the **16**. To understand this, recall that the  $\Gamma$ -matrices of the 10d Clifford algebra give rise to five independent, anticommuting ‘creation-annihilation’ operators  $\Gamma^{a\pm} = (\Gamma^{2a-1} \pm i\Gamma^{2a})/2$  with  $a = 1, \dots, 5$ . These correspond to five fermionic harmonic oscillators or “spin”  $1/2$  systems. The 32-dimensional tensor product of those is reducible since the 10d rotation generators  $M_{mn} = -i[\Gamma^m, \Gamma^n]/4$  ( $m, n = 1, \dots, 10$ ) always flip an even number of “spins”. This gives rise to the **16** as displayed in Table 16.1.

Next, one also recalls that the natural embedding of  $SU(5)$  in  $SO(10)$  relies on ‘pairing up’ real dimensions,  $R^{10} \equiv C^5$ , similarly to the pairing up of  $\Gamma^m$ s used above. This makes it clear how to associate one  $|\pm\rangle$  system to each complex dimension of  $SU(5)$ , which explains the labeling of the “spin” columns in Table 16.1: The first three and last two “spins” correspond to  $SU(3)_C$  and  $SU(2)_L$  respectively. In fact, an  $SU(3)_C$  rotation just raises one color index and lowers another, changing colors  $\{r, b, y\}$ , or changes relative phases between the three spin states. Similarly, an  $SU(2)_L$  rotation raises one weak index and lowers another, thereby flipping the weak isospin from up to down or vice versa, or changes the relative phase between the two spin states. In this representation  $U(1)_Y$  hypercharge is simply given by  $Y = -2/3(\sum \text{color spins}) + (\sum \text{weak spins})$ .  $SU(5)$  rotations corresponding to  $X$  bosons then raise (or lower) a color index, while at the same time lowering (or raising) a weak index. It is easy to see that such rotations can mix the states  $\{Q, u^c, e^c\}$  and  $\{d^c, L\}$  among themselves and  $\nu^c$  is a singlet. Since  $SO(10)$  has 45 generators, additional 21 gauge bosons are introduced including the  $U(1)_X$  above. The 20 new  $SO(10)$  rotations not in  $SU(5)$  are then given by either raising any two spins or lowering them. With these rotations, **1** and **5** are connected with **10**. The last  $SO(10)$  rotation changes phases of states with weight  $2(\sum \text{color spins}) + 2(\sum \text{weak spins})$ , which corresponds to  $U(1)_X$ .

**Table 16.1:** Quantum numbers of **16**-dimensional representation of  $SO(10)$ .

state	Y	Color	Weak	$ SU(5) $	$ SO(10) $
$\nu^c$	0	---	--	<b>1</b>	<b>16</b>
$e^c$	2	---	++		
$u_r$	1/3	+--	-+	<b>10</b>	
$d_r$	1/3	+--	+-		
$u_b$	1/3	-+-	-+	<b>10</b>	
$d_b$	1/3	-+-	+-		
$u_y$	1/3	--+	-+		
$u_y$	1/3	--+	+-		
$u_r^c$	-4/3	-++	--	<b>5</b>	
$u_b^c$	-4/3	-++	--		
$u_y^c$	-4/3	++-	--		
$d_r^c$	2/3	-++	++	<b>5</b>	
$d_b^c$	2/3	+-+	++		
$d_y^c$	2/3	++-	++		
$\nu$	-1	+++	-+		
$e$	-1	+++	+-		

$SO(10)$  has two inequivalent maximal subgroups and hence breaking patterns,  $SO(10) \rightarrow SU(5) \times U(1)_X$  and  $SO(10) \rightarrow SU(4)_C \times SU(2)_L \times SU(2)_R$ . In the first case, one can carry on breaking to  $G_{SM} \subset SU(5)$  precisely as in the minimal  $SU(5)$  case above. Alternatively, one can identify  $U(1)_Y$  as an appropriate linear

combination of  $U(1)_X$  and the  $U(1)$  factor from  $SU(5)$ , leading to the so-called flipped  $SU(5)$  [16] as an intermediate step in breaking  $SO(10)$  to  $G_{SM}$ . In the second case, we have an intermediate Pati-Salam model thanks to the branching rule **16** = **(4, 2, 1)** + **( $\bar{4}$ , 1, 2)**. Finally,  $SO(10)$  can break directly to the SM at  $M_G$ . Gauge coupling unification remains intact in the case of this ‘direct’ breaking and for the breaking pattern  $SO(10) \rightarrow SU(5) \rightarrow G_{SM}$  (with  $SU(5)$  broken at  $M_G$ ). In the case of intermediate-scale Pati-Salam or flipped  $SU(5)$  models, gauge coupling predictions are modified. The Higgs multiplets in minimal  $SO(10)$  come from the fundamental representation, **10<sub>H</sub>** = **5<sub>H</sub>** +  **$\bar{5}_H$** . Note, only in  $SO(10)$  does the representation type distinguish SM matter from Higgs fields.

Finally, larger symmetry groups can be considered. For example, the exceptional group  $E_6$  has maximal subgroup  $SO(10) \times U(1)$  [17]. Its fundamental representation branches as **27** = **16<sub>1</sub>** + **10<sub>-2</sub>** + **1<sub>4</sub>**. Another maximal subgroup is  $SU(3)_C \times SU(3)_L \times SU(3)_R \subset E_6$  with branching rule **27** = **(3, 3, 1)** + **( $\bar{3}$ , 1,  $\bar{3}$ )** + **(1,  $\bar{3}$ , 3)**. Independently of any underlying  $E_6$ , the group  $[SU(3)]^3$  with additional permutation symmetry  $Z_3$  interchanging the three factors can be considered. This is known as “trification” [18]. The  $E_6 \rightarrow [SU(3)]^3$  breaking pattern has been used in phenomenological analyses of the heterotic string [19]. However, in larger symmetry groups, such as  $E_6$ ,  $SU(6)$ , etc., there are now many more states which have not been observed and must be removed from the effective low-energy theory.

Intriguingly, the logic by which  $G_{SM}$  is a maximal subgroup of  $SU(5)$ , which together with  $U(1)_X$  is a maximal subgroup of  $SO(10)$ , continues in a very elegant and systematic way up to the largest exceptional group. The resulting famous breaking chain  $E_8 \rightarrow E_7 \rightarrow E_6 \rightarrow SO(10) \rightarrow SU(5) \rightarrow G_{SM}$  together with the special role played by  $E_8$  in group and in string theory is a tantalizing hint at deeper structures. However, since all representations of  $E_8$  and  $E_7$  are real and can not lead to 4d chiral fermions, this is necessarily outside the 4d GUT framework.

### 16.3. GUT breaking and doublet-triplet splitting

In the standard, 4d field-theoretic approach to GUTs, the unified gauge group is broken spontaneously by an appropriate GUT Higgs sector. Scalar potentials (or superpotentials in SUSY GUTs) exist whose vacua spontaneously break  $SU(5)$  or  $SO(10)$ . While these potentials are ad hoc (just like the Higgs potential in the SM), the most naive expectation is that all their dimensionful parameters are  $O(M_G)$ . In the simplest case of  $SU(5)$ , the **24** (adjoint) GUT Higgs develops a VEV along the  $G_{SM}$ -singlet direction as  $\langle \Phi \rangle \propto \text{diag}(-2/3, -2/3, -2/3, 1, 1)$ . In order for  $SO(10)$  to break to  $SU(5)$ , the **16** or **126**, which have a  $G_{SM}$ -singlet with non-zero  $U(1)_X$  charge, get a VEV.

The masses of doublet and triplet in the **5<sub>H</sub>** (and  **$\bar{5}_H$** ) generically split due to their coupling to the GUT Higgs. In addition, both the doublet and the triplet mass also get an equal contribution from an  $SU(5)$ -invariant GUT-scale mass term. Without any further structure, an extreme fine-tuning between two large effects is then necessary to keep the doublet mass at the electroweak scale. Supersymmetry plays an important role in forbidding large radiative correction to the doublet mass due to the non-renormalization theorem [7]. However, even in this case we have to fine tune parameters at tree level. This is the doublet-triplet splitting problem which, in the SUSY context, is clearly related the  $\mu$ -term problem of the MSSM (the smallness of the coefficient of  $\mu H_u H_d$ ).

Several mechanisms for natural doublet-triplet splitting have been suggested under the assumption of supersymmetry, such as the sliding singlet [20], missing partner [21] or missing VEV [22], and pseudo-Nambu-Goldstone boson mechanisms [23]. Particular examples of the missing partner mechanism for  $SU(5)$  [24], the missing VEV mechanism for  $SO(10)$  [25,26] and the pseudo-Nambu-Goldstone boson mechanism for  $SU(6)$  [23,27] have been shown to be consistent with gauge coupling unification and nucleon decay. From the GUT-scale perspective, one is satisfied if the triplets are naturally heavy and the doublets are massless ( $\mu \simeq 0$ ). There are also several mechanisms for resolving the subsequent issue of why  $\mu$  is of order

the SUSY breaking scale [28]. \* For a review of the  $\mu$  problem and some suggested solutions in SUSY GUTs and string theory, see [29,30,31,32] and references therein.

In general, GUT-breaking sectors successfully resolving the doublet-triplet splitting problem, dynamically stabilizing all GUT-scale VEVs and allowing for realistic neutrino masses and Yukawa couplings (including the GUT-symmetry violation in the latter) require a number of ingredients. However, for validity of the effective theory, introduction of higher or many representations is limited, otherwise a Landau pole may appear below the Planck scale. In addition, GUTs are only effective theories below the Planck scale in the 4d field-theoretic approach. Since  $M_G$  is close to this scale, the effects of higher-dimension operators are not obviously negligible. In particular, operators including the GUT-breaking Higgs may affect low-energy predictions, such as quark and lepton masses.

Thus, especially in the context of GUT breaking and doublet-triplet splitting, models beyond 4d field theory appear attractive. While this is mainly the subject of the next section, some advantages can already be noted: In models with extra dimensions, in particular string constructions, GUT breaking may occur due to boundary conditions in the compactified dimensions [33,34,35,36]. No complicated GUT breaking sector is then required. Moreover, boundary conditions can give mass only to the triplet, leaving the doublet massless. This is similar to the ‘missing partner mechanism’ since the effective mass term does not ‘pair up’ the triplets from  $\mathbf{5}_H$  and  $\bar{\mathbf{5}}_H$  but rather each of them with further fields which are automatically present in the higher-dimensional theory. This can eliminate dimension-five nucleon decay (cf. Sec. 16.6).

## 16.4. String-theoretic and Higher-dimensional Unified Models

As noted earlier, the GUT scale is dangerously close to the scale of quantum gravity. It may hence be necessary to discuss unified models of particle physics in the latter, more ambitious context. Among the models of quantum gravity, superstring or M-theory stands out as the best-studied and technically most developed proposal, possessing in particular a high level internal, mathematical consistency. For our purposes, it is sufficient to know that five 10d and one 11d low-energy effective supergravity theories arise in this setting (cf. [37] and refs. therein).

Grand unification is realized most naturally in the context of the two ‘heterotic’ theories with gauge groups  $E_8 \times E_8$  and  $SO(32)$  respectively [35] (see [38] for some of the more recent results). Justified in part by the intriguing breaking path  $E_8 \rightarrow \dots \rightarrow G_{SM}$  mentioned above, the focus has historically largely been on  $E_8 \times E_8$ . To describe particle physics, solutions of the 10d theory with geometry  $R^{1,3} \times M_6$  are considered, where  $M_6$  is a Calabi-Yau (CY) 3-fold (with 6 real dimensions). The background solution involves expectation values of higher-dimensional components of the  $E_8 \times E_8$  gauge fields. This includes both Wilson lines [33] and non-vanishing field-strength and leads, in general, to a reduced gauge symmetry and to chirality in the resulting 4d effective theory. The 4d fermions arise from 10d gauginos.

Given an appropriate embedding of  $G_{SM}$  in  $E_8 \times E_8$ , gauge coupling unification is automatic at leading order. Corrections arise mainly through (string)-loop effects and are similar to the familiar field-theory thresholds of 4d GUTs [39]. Thus, one may say that coupling unification is a generic prediction in spite of the complete absence\* of a 4d GUT at any energy scale. This absence is both an advantage and a weakness. On the up side, GUT breaking and doublet-triplet splitting [41] are more naturally realized and dimension-five nucleon decay is relatively easy to avoid. On the down side, there is no reason

to expect full GUT representations in the matter sector and flavor model building is much less tied to the GUT structure than in 4d.

One technical problem of heterotic constructions is the dependence on the numerous size and shape parameters of  $M_6$  (the so-called moduli), the stabilization of which is poorly understood (see [42] for recent developments). Another is the sheer mathematical complexity of the analysis, involving in particular the study of (non-Abelian) gauge-bundles on CY spaces [43] (see however [44]).

An interesting sub-chapter of heterotic string constructions is represented by orbifold models [34]. Here the internal space is given by a six-torus, modded out by a discrete symmetry group (e.g.  $T^6/Z_n$ ). More recent progress is reported in [45], including in particular the systematic exploration of the phenomenological advantages of so-called ‘non-prime’ (referring to  $n$ ) orbifolds. The symmetry breaking to  $G_{SM}$  as well as the survival of Higgs doublets without triplet partners is ensured by the appropriate embedding of the discrete orbifold group in  $E_8 \times E_8$ .

String theory on such spaces, which are locally flat but include singularities, is much more calculable than in the CY case. The orbifold geometries can be viewed as singular limits of CYs.

An even simpler approach to unified models, which includes many of the advantages of full-fledged string constructions, is provided by Orbifold GUTs [36]. These are (mostly) 5d or 6d SUSY field theories with unified gauge group (e.g.  $SU(5)$  or  $SO(10)$ ), broken in the process of compactifying to 4d. To give a particularly simple example, consider  $SU(5)$  on  $R^{1,3} \times S^1/(Z_2 \times Z'_2)$ . Here the compact space is an interval and the embedding of  $Z'_2$  in the hypercharge direction of  $SU(5)$  realizes the breaking to  $G_{SM}$ . Concretely, 5d  $X$  bosons are given Neumann BCs at one endpoint of the interval and thus have no Kaluza-Klein (KK) zero mode. Their lightest modes have mass  $\sim 1/R$ , making the KK-scale the effective GUT scale. As an implication, the boundary theory has no  $SU(5)$  invariance. Nevertheless, since the  $SU(5)$ -symmetric 5d bulk dominates 4d gauge couplings, unification remains a prediction. Many other features but also problems of 4d GUTs can be circumvented, especially doublet-triplet splitting is easily realized.

With the advent of the string-theory ‘flux landscape’ [46], which is best understood in 10d type-IIB supergravity, the focus in string model building has shifted to this framework. While type II string theories have no gauge group in 10d, brane-stacks support gauge dynamics. A particularly appealing setting (see e.g. [47]) is provided by type IIB models with D7 branes (defining 8d submanifolds). However, in the  $SO(10)$  context the  $\mathbf{16}$  is not available and, for  $SU(5)$ , the top-Yukawa coupling vanishes at leading order [48]. As a crucial insight, this can be overcome on the non-perturbative branch of type IIB, also known as F-theory [49,50]. This setting allows for more general branes, thus avoiding constraints of the  $Dp$ -brane framework. GUT breaking can be realized using hypercharge flux (the VEV of the  $U(1)_Y$  field strength), an option not available in heterotic models. The whole framework combines the advantages of the heterotic or higher-dimensional unification approach with the more recent progress in understanding moduli stabilization. It thus represents at this moment the most active and promising branch of theory-driven GUT model building (see e.g. [51] and refs. therein).

As a result of the flux-breaking, a characteristic ‘type IIB’ or ‘F-theoretic’ tree-level correction to gauge unification arises [52]. The fact that this correction can be rather significant numerically is occasionally held against the framework of F-theory GUTs. However, at a parametric level, this correction nevertheless behaves like a 4d threshold, i.e., it provides  $\mathcal{O}(1)$  additive contributions to the inverse 4d gauge coupling  $\alpha_i^{-1}(M_{GUT})$ .

A final important issue in string GUTs is the so-called string-scale/GUT-scale problem [53]. It arises since, in heterotic compactifications, the Planck scale and the high-scale value of the gauge coupling unambiguously fix the string-scale to about  $10^{18}$  GeV. As the compactification radius  $R$  is raised above the string length, the GUT scale (identified with  $1/R$ ) goes down and the string coupling goes up. Within the domain of perturbative string theory, a gap of about a factor  $\sim 20$  remains between the lowest GUT scale achievable in this way and the phenomenological goal of  $2 \times 10^{16}$  GeV. The situation

\* The solution of [28] relies on the absence of the fundamental superpotential term  $\mu H_u H_d$  (or  $\mu \mathbf{5}_H \bar{\mathbf{5}}_H$ ). This is ensured by a  $U(1)_R$ . The latter clashes with typical superpotentials for the GUT breaking sector. However, higher-dimensional or stringy GUTs, where the triplet Higgs is simply projected out, can be consistent with the  $U(1)_R$  symmetry.

\* See however [40].

can be improved by venturing into the non-perturbative regime [53] or by considering ‘anisotropic’ geometries with hierarchically different radii  $R$  [53,54].

In F-theory GUTs, the situation is dramatically improved since the gauge theory lives only in four out of the six compact dimensions. This allows for models with a ‘decoupling limit’, where the GUT scale is parametrically below the Planck scale [50]. However, moduli stabilization may not be without problems in such constructions, in part due to a tension between the required large volume and the desirable low SUSY breaking scale.

### 16.5. Gauge Coupling Unification

The quantitative unification of the three SM gauge couplings at the energy scale  $M_G$  is one of the cornerstones of the GUT paradigm. It is obviously of direct phenomenological relevance. Gauge coupling unification is best understood in the framework of effective field theory (EFT) [55]. In the simplest case, the relevant EFT at energies  $\mu \gg M_G$  has a unified gauge symmetry (say  $SU(5)$  for definiteness) and a single running gauge coupling  $\alpha_G(\mu)$ . At energies  $\mu \ll M_G$ , states with mass  $\sim M_G$  (such as  $X$  bosons, GUT Higgs, color-triplet Higgs) have to be integrated out. The EFT now has three independent couplings and SM (or MSSM) matter content. One-loop renormalization group equations readily allow for an extrapolation to the weak scale,

$$\alpha_i^{-1}(m_Z) = \alpha_G^{-1}(M_G) + \frac{b_i}{2\pi} \log \left( \frac{M_G}{m_Z} \right) + \delta_i. \quad (16.3)$$

Here we defined  $\delta_i$  to absorb all sub-leading effects, including threshold corrections at or near the weak scale (e.g. from superpartners and the additional Higgs bosons in the case of SUSY). We will discuss them momentarily.

It is apparent from Eq. (16.3) that the three low-scale couplings can be very different. This is due to the large energy range  $m_Z \ll \mu \ll M_G$  and the non-universal  $\beta$ -function coefficients ( $b_i^{\text{SM}} = \{41/10, -19/6, -7\}$  or  $b_i^{\text{MSSM}} = \{33/5, 1, -3\}$ ). Incomplete GUT multiplets, such as gauge and Higgs bosons in the SM and also their superpartners and the additional Higgs bosons in the MSSM, contribute to the differences between the  $\beta$  functions. Inverting the argument, one expects that extrapolating the measured couplings to the high scale, we find quantitative unification at  $\mu \sim M_G$ . While this fails in the SM, it works intriguingly well in the MSSM (cf. Fig. 1).

The three equations contained in (Eq. (16.3)) can be used to determine the three ‘unknowns’  $\alpha_3(m_Z)$ ,  $\alpha_G(M_G)$  and  $M_G$ , assuming that all other parameters entering the equations are given. Focusing on the SUSY case and using the  $\overline{\text{MS}}$  coupling constants  $\alpha_{\text{EM}}^{-1}(m_Z)$  and  $\sin^2 \theta_W(m_Z)$  from [56],

$$\alpha_{\text{EM}}^{-1}(m_Z) = 127.940 \pm 0.014, \quad (16.4)$$

$$\sin^2 \theta_W(m_Z) = 0.23126 \pm 0.00005, \quad (16.5)$$

as input, one determines  $\alpha_{1,2}^{-1}(m_Z)$ , which then gives

$$\alpha_G^{-1}(M_G) \simeq 24.3 \quad \text{and} \quad M_G \simeq 2 \times 10^{16} \text{ GeV}. \quad (16.6)$$

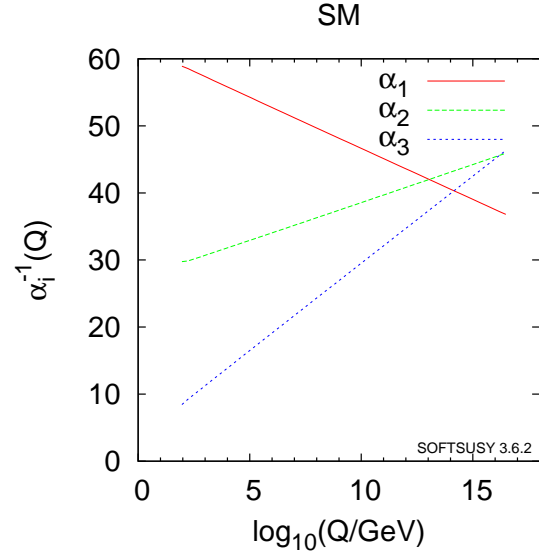
Here we have set  $\delta_i = 0$  for simplicity. Crucially, one in addition obtains a prediction for the low-energy observable  $\alpha_3$ ,

$$\alpha_3^{-1}(m_Z) = -\frac{5}{7} \alpha_1^{-1}(m_Z) + \frac{12}{7} \alpha_2^{-1}(m_Z) + \Delta_3, \quad (16.7)$$

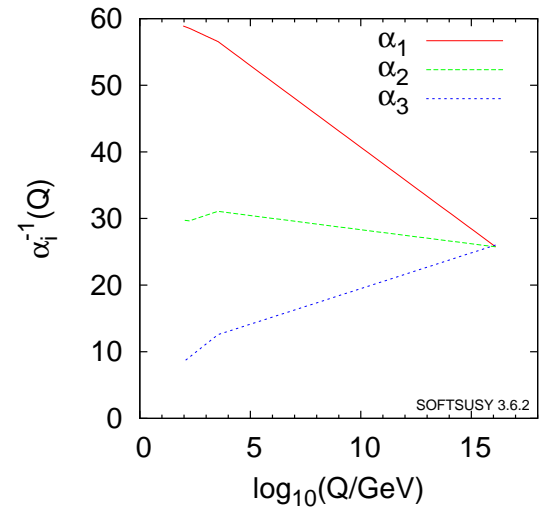
where

$$\Delta_3 = \frac{5}{7} \delta_1 - \frac{12}{7} \delta_2 + \delta_3. \quad (16.8)$$

Here we followed the elegant formulation in Ref. [57] of the classical analyses of [4]. Of course, it is a matter of convention which of the three low-energy gauge coupling parameters one ‘predicts’ and indeed, early work on the subject discusses the prediction of  $\sin^2 \theta_W$  in terms of  $\alpha_{\text{EM}}$  and  $\alpha_3$  [58,59].



MSSM:  $m_0 = M_{1/2} = 2 \text{ TeV}$ ,  $A_0 = 0$ ,  $\tan \beta = 30$



**Figure 16.1:** Running couplings in SM and MSSM using two-loop RG evolution. The SUSY threshold at 2 TeV is clearly visible on the r.h. side. (We thank Ben Allanach for providing the plots created using SOFTSUSY [61].)

Remarkably, the leading order result (i.e. Eq. (16.7) with  $\delta_i = 0$ ) is in excellent agreement with experiments [56]:

$$\alpha_3^{\text{LO}}(m_Z) = 0.117 \quad \text{vs.} \quad \alpha_3^{\text{EXP}}(m_Z) = 0.1185 \pm 0.0006. \quad (16.9)$$

However, this near perfection is to some extent accidental. To see this, we now discuss the various contributions to the  $\delta_i$  (and hence to  $\Delta_3$ ).

The two-loop running correction from the gauge sector  $\Delta_3^{(2)}$  and the low-scale threshold correction  $\Delta_3^{(l)}$  from superpartners can be summarized as [57]

$$\Delta_3^{(2)} \simeq -0.82 \quad \text{and} \quad \Delta_3^{(l)} \simeq \frac{19}{28\pi} \log \left( \frac{m_{\text{SUSY}}}{m_Z} \right). \quad (16.10)$$

The relevant scale  $m_{\text{SUSY}}$  can be estimated as [60]

$$m_{\text{SUSY}} \rightarrow m_H^{3/19} m_{\tilde{H}}^{12/19} m_{\tilde{W}}^{4/19} \times \left( \frac{m_{\tilde{W}}}{m_{\tilde{g}}} \right)^{28/19} \left( \frac{m_{\tilde{t}}}{m_{\tilde{q}}} \right)^{3/19}, \quad (16.11)$$

where  $m_H$  stands for the masses of non-SM Higgs states and superpartner masses are given in self-evident notation. Detailed

analyses including the above effects are best done using appropriate software packages, such as SOFTSUSY [61] (or alternatively SuSpect or SPheno [62]). See also [61] for references to the underlying theoretical two-loop analyses.

To get a very rough feeling for these effects, let us assume that all superpartners are degenerate at  $m_{\text{SUSY}} = 1$  TeV, except for heavier gluinos:  $m_{\tilde{W}}/m_{\tilde{g}} \simeq 1/3$ . This gives  $\Delta_3^{(l)} \simeq -0.35 + 0.22 \ln(m_{\text{SUSY}}/m_Z) \simeq 0.18$ . The resulting prediction of  $\alpha_3(m_Z) \simeq 0.126$  significantly upsets the perfect one-loop agreement found earlier. Before discussing this issue further, it is useful to introduce yet another important type of correction, the high- or GUT-scale thresholds.

To discuss high-scale thresholds, let us set all other corrections to zero for the moment and write down a version of Eq. (16.3) that captures the running near and above the GUT scale more correctly. The threshold correction at one-loop level can be evaluated accurately by the simple step-function approximation for the  $\beta$  functions in the  $\overline{\text{DR}}$  scheme\* [66],

$$\alpha_i^{-1}(m_Z) = \alpha_i^{-1}(\mu) + \frac{1}{2\pi} \left[ b_i \ln \frac{\mu}{m_Z} + b_i^C \ln \frac{\mu}{M_C} + b_i^X \ln \frac{\mu}{M_X} + b_i^\Phi \ln \frac{\mu}{M_\Phi} \right]. \quad (16.12)$$

Here we started the running at some scale  $\mu \gg M_G$ , including the contribution of the minimal set of states relevant for the transition from the high-scale  $SU(5)$  model to the MSSM. These are the color-triplet Higgs multiplets with mass  $M_C$ , massive vector multiplets of  $X$ -bosons with mass  $M_X$  (including GUT Higgs degrees of freedom), and the remaining GUT-Higgs fields and superpartners with mass  $M_\Phi$ . The coefficients  $b_i^{C,X,\Phi}$  can be found in Ref. [67]. Crucially, the  $b_i$  in Eq. (16.12) conspire to make the running GUT-universal at high scales, such that the resulting prediction for  $\alpha_3$  does not depend on the value of  $\mu$ .

To relate this to our previous discussion, we can, for example, define  $M_G \equiv M_X$  and then choose  $\mu = M_G$  in Eq. (16.12). This gives the high-scale threshold corrections

$$\delta_i^{(h)} = \frac{1}{2\pi} \left[ b_i^C \ln \frac{M_G}{M_C} + b_i^\Phi \ln \frac{M_G}{M_\Phi} \right], \quad (16.13)$$

and a corresponding correction  $\Delta_3^{(h)}$ . To get some intuition for the magnitude, one can furthermore assume  $m_\Phi = M_G$ , finding (with  $b_i^C = \{2/5, 0, 1\}$ )

$$\Delta_3^{(h)} = \frac{9}{14\pi} \ln \left( \frac{M_G}{M_C} \right). \quad (16.14)$$

To obtain the desired effect of +0.64, the triplet Higgs would have to be by about a factor 20 lighter than the GUT scale. While this is ruled out by nucleon decay in the minimal model [68] as will be discussed Sec. 16.6, it is also clear that threshold corrections of this order of magnitude can, in general, be realized with a certain amount of GUT-scale model building, e.g. in specific  $SU(5)$  [24] or  $SO(10)$  [25,26] constructions. It is, however, a significant constraint on the 4d GUT sector of the theory.

The above analysis implicitly assumes universal soft SUSY breaking masses at the GUT scale, which directly affect the spectrum of SUSY particles at the weak scale. In the simplest case we have a universal gaugino mass  $M_{1/2}$ , a universal mass for squarks and sleptons  $m_{16}$  and a universal Higgs mass  $m_{10}$ , as motivated by  $SO(10)$ . In some cases, threshold corrections to gauge coupling unification can be exchanged for threshold corrections to soft SUSY parameters (see [69] and refs. therein). For example, if gaugino masses were not unified at  $M_G$  and, in particular, gluinos were lighter than winos at the weak scale (cf. Eq. (16.11)), then it is possible that, due to weak scale threshold corrections, a much smaller or even slightly positive

threshold correction at the GUT scale would be consistent with gauge coupling unification [70].

It is also noteworthy that perfect unification can be realized without significant GUT-scale corrections, simply by slightly raising the (universal) SUSY breaking scale. In this case the dark matter abundance produced by thermal processes in the early universe (if the lightest neutralino is the dark matter particle) is too high. However, even if the gaugino mass in the MSSM is about 1 TeV, if the Higgsino and the non-SM Higgs boson masses are about 10-100 TeV, the effective SUSY scale can be raised [71]. This setup is realized in split SUSY [72] or the pure gravity mediation model [73] based on anomaly mediation [74]. Since the squarks and sleptons are much heavier than the gaugino masses in those setups, a gauge hierarchy problem is reintroduced. The facts that no superpartners have so far been seen at the LHC and that the observed Higgs mass favors heavier stop masses than about 1 TeV force one to accept a certain amount of fine-tuning anyway.

For non-SUSY GUTs or GUTs with a very high SUSY breaking scale to fit the data, new light states in incomplete GUT multiplets or multiple GUT breaking scales are required. For example, non-SUSY models  $SO(10) \rightarrow SU(4)_C \times SU(2)_L \times SU(2)_R \rightarrow \text{SM}$ , with the second breaking scale of order an intermediate scale, determined by light neutrino masses using the see-saw mechanism, can fit the low-energy data for gauge couplings [75] and at the same time survive nucleon decay bounds [76]. Alternatively, one can appeal to string-theoretic corrections discussed in Sec. 16.4 to compensate for a high SUSY breaking scale. This has, for example, been concretely analyzed in the context of F-theory GUTs in [77].

In 5d or 6d orbifold GUTs, certain “GUT scale” threshold corrections come from the Kaluza-Klein modes between the compactification scale,  $M_c$ , and the effective cutoff scale  $M_*$ . In string theory, this cutoff scale is the string scale. Gauge coupling unification at two loops then constrains the values of  $M_c$  and  $M_*$ . Typically, one finds  $M_c$  to be lower than the 4d GUT scale. Since the  $X$ -bosons, responsible for nucleon decay, get mass at the compactification scale, this has significant consequences for nucleon decay.

Finally, it has been shown that non-supersymmetric GUTs in warped 5d orbifolds can be consistent with gauge coupling unification. This assumes (in 4d language) that the r.h. top quark and the Higgs doublets are composite-like objects with a compositeness scale in the TeV range [79].

## 16.6. Nucleon Decay

Quarks and leptons are indistinguishable in any 4d GUT, and both the baryon ( $B$ ) and lepton number ( $L$ ) are not conserved. This leads to baryon-number-violating nucleon decay. In addition to baryon-number violation, lepton-number violation is also required for nucleon decay since, in the SM, leptons are the only free fermions which are lighter than nucleons. The lowest-dimension operators relevant for nucleon decay are  $(B+L)$  violating dimension-six four-fermion-terms since all baryon-violating operators with dimension less than seven preserve  $(B-L)$  [80]. In  $SU(5)$  GUTs, they are induced by  $X$  boson exchange. These operators are suppressed by  $(1/M_G^2)$ , and the nucleon lifetime is given by  $\tau_N \propto M_G^4/(\alpha_G^2 m_p^5)$  ( $m_p$  is proton mass). The dominant decay mode of the proton (and the baryon-violating decay mode of the neutron), via  $X$  boson exchange, is  $p \rightarrow e^+ \pi^0$  ( $n \rightarrow e^+ \pi^-$ ). In any simple gauge symmetry, with one universal GUT coupling  $\alpha_G$  and scale  $M_G$ , the nucleon lifetime from gauge boson exchange is calculable. Hence, the GUT scale may be directly observed via the extremely rare decay of the nucleon. Experimental searches for nucleon decay began with the Kolar Gold Mine, Homestake, Soudan, NUSEX, Frejus, HPW, and IMB detectors [58]. The present experimental bounds come from Super-Kamiokande and Soudan II. We discuss these results shortly. While non-SUSY GUTs are constrained by the non-observation of nucleon decay, a precise and general statement is hard to make. The reason is that gauge couplings do not unify with

\* The  $\overline{\text{DR}}$  scheme is frequently used in a supersymmetric regularization [63]. The renormalization transformation of the gauge coupling constants from  $\overline{\text{MS}}$  to  $\overline{\text{DR}}$  scheme is given in Ref. [64]. For an alternative treatment using holomorphic gauge couplings and NSVZ  $\beta$ -functions see e.g. [65].

\* It is interesting to note that a ratio  $M_*/M_c \sim 100$ , needed for gauge coupling unification to work in orbifold GUTs, is typically the maximum value for this ratio consistent with perturbativity [78].

just the SM particle content. Once extra states or large thresholds are included to ensure precision unification, a certain range of unification scales is allowed. By contrast, in SUSY GUTs one generically has  $M_G \sim 2 \times 10^{16}$  GeV. Hence dimension-six baryon-number-violating operators are predicted to induce a lifetime of about  $\tau_p \sim 10^{36}$  years.

However, in SUSY GUTs there are additional sources for baryon and/or lepton-number violation – dimension-four and five operators [13]. These arise since, in the SUSY SM, quarks and leptons have scalar partners (squarks and sleptons). Although our notation does not change, when discussing SUSY models our fields are chiral superfields and both fermionic and bosonic matter is implicitly represented by those. In this language, baryon- and/or lepton-number-violating dimension-four and five operators are given as so-called  $F$  terms of products of chiral superfields, which contain two fermionic components and the rest scalars or products of scalars. Within the context of  $SU(5)$  the dimension-four and five operators have the form

$$(10 \bar{5} \bar{5}) \supset (u^c d^c d^c) + (Q L d^c) + (e^c L L),$$

$$(10 10 10 \bar{5}) \supset (Q Q Q L) + (u^c u^c d^c e^c)$$

+  $B$ - and  $L$ -conserving terms,

respectively. The dimension-four operators are renormalizable, with dimensionless couplings similar to Yukawa couplings. By contrast, the dimension-five operators have a dimensionful coupling of order  $(1/M_G)$ . They are generated by integrating out the color-triplet Higgs with GUT-scale mass. Note that both triplet Higgsinos (due to their fermionic nature) and Higgs scalars (due to their mass-enhanced trilinear coupling with matter) contribute to the operators.

The dimension-four operators violate either baryon number or lepton number. The nucleon lifetime is extremely short if both types of dimension-four operators are present in the SUSY SM since squark or slepton exchange induces the dangerous dimension-six SM operators. Even in the case that they violate baryon number or lepton number only but not both, they are constrained by various phenomena [81]. For example, the primordial baryon number in the universe is washed out unless the dimensionless coupling constants are less than  $10^{-7}$ . Both types of operators can be eliminated by requiring  $R$  parity, which distinguishes Higgs from ordinary matter multiplets.  $R$  parity [82] or its cousin, matter parity [12,83], act as  $F \rightarrow -F$ ,  $H \rightarrow H$  with  $F = \{10, \bar{5}\}$ ,  $H = \{\bar{5}_H, 5_H\}$  in  $SU(5)$ . This forbids the dimension-four operator  $(10 \bar{5} \bar{5})$ , but allows the Yukawa couplings for quark and lepton masses of the form  $(10 \bar{5} \bar{5}_H)$  and  $(10 10 5_H)$ . It also forbids the dimension-three, lepton-number-violating operator  $(\bar{5} 5_H) \supset (L H_u)$  as well as the dimension-five, baryon-number-violating operator  $(10 10 10 \bar{5}_H) \supset (Q Q Q H_d) + \dots$ . In  $SU(5)$ , the Higgs multiplet  $\bar{5}_H$  and the matter multiplets  $\bar{5}$  have identical gauge quantum numbers. In  $E_6$ , Higgs and matter multiplets could be unified within the fundamental  $27$  representation. Only in  $SO(10)$  are Higgs and matter multiplets distinguished by their gauge quantum numbers. Moreover the  $Z_4$  center of  $SO(10)$  distinguishes  $10_s$  from  $16_s$  and can be associated with  $R$  parity [84].

The dimension-five baryon-number-violating operators may also be forbidden at tree level by certain symmetries consistent with  $SU(5)$  [13]. However, these symmetries are typically broken by the VEVs responsible for the color-triplet Higgs masses. Consequently the dimension-five operators are generically generated via the triplet Higgs exchange in SUSY  $SU(5)$  GUTs, as mentioned above. Hence, the triplet partners of Higgs doublets must necessarily obtain mass of order the GUT scale. In addition, it is also important to note that Planck or string scale physics may independently generate the dimension-five operators, even without a GUT. These contributions must be suppressed by some underlying symmetry; for example, the same flavor symmetry which may be responsible for hierarchical fermion Yukawa matrices.

As a general remark, appealing to global symmetries to suppress specific interactions may not always be as straightforward as it naively seems. Indeed, there are two possibilities: On the one hand, the relevant symmetry might be gauged at a higher scale. Effects of the VEVs responsible for the spontaneous breaking are then in principle dangerous and need to be quantified.

On the other hand, the symmetry might be truly only global. This must e.g. be the case for anomalous symmetries, which are then also violated by field-theoretic non-perturbative effects. The latter can in principle be exponentially small. It is, however, widely believed that global symmetries are always broken in quantum gravity (see e.g. [85]). One then needs to understand which power or functional form the Planck scale suppression of the relevant interaction has. For example, dimension-five baryon number violating operators suppressed by just one unit of the Planck or string scale are completely excluded.

In view of the above, it is also useful to recall that in string models 4d global symmetries generally originate in higher-dimensional gauge symmetries. Here ‘global’ implies that the gauge boson has acquired a Stückelberg-mass. This is a necessity in the anomalous case (Green-Schwarz mechanism) but can also happen to non-anomalous symmetries. One expects no symmetry violation beyond the well-understood non-perturbative effects. Discrete symmetries arise as subgroups of continuous gauge symmetries, such as  $Z_N \subset U(1)$ . In particular, non-anomalous subgroups of Stückelberg-massive  $U(1)$ s represent unbroken discrete gauge symmetries and as such are non-perturbatively exact (see e.g. [86]). Of course, such discrete gauge symmetries may also arise as remnants of continuous gauge symmetries after conventional 4d spontaneous breaking.

Dimension-five operators include squarks and/or sleptons. To allow for nucleon decay, these must be converted to light quarks or leptons by exchange of a gaugino or Higgsino in the SUSY SM. The nucleon lifetime is proportional to  $M_G^2 m_{\text{SUSY}}^2 / m_p^5$ , where  $m_{\text{SUSY}}$  is the SUSY breaking scale. Thus, dimension-five operators may predict a shorter nucleon lifetime than dimension-six operators. Unless accidental cancellations are present, the dominant decay modes from dimension-five operators include a  $K$  meson, such as  $p \rightarrow K^+ \bar{\nu}$  ( $n \rightarrow K^0 \bar{\nu}$ ). This is due to a simple symmetry argument: The operators are given as  $(Q_i Q_j Q_k L_l)$  and  $(u_i^c u_j^c d_k^c e_l^c)$ , where  $i, j, k, l$  ( $= 1-3$ ) are family indices and color and weak indices are implicit. They must be invariant under  $SU(3)_C$  and  $SU(2)_L$  so that their color and weak doublet indices must be anti-symmetrized. Since these operators are given by bosonic superfields, they must be totally symmetric under interchange of all indices. Thus the first operator vanishes for  $i = j = k$  and the second vanishes for  $i = j$ . Hence a second or third generation member exists in the dominant modes of nucleon decay unless these modes are accidentally suppressed [83].

Recent Super-Kamiokande bounds on the proton lifetime severely constrain the dimension-six and five operators. With 306 kton-years of data they find  $\tau_p / \text{Br}(p \rightarrow e^+ \pi^0) > 1.67 \times 10^{34}$  years and  $\tau_p / \text{Br}(p \rightarrow K^+ \bar{\nu}) > 6.6 \times 10^{33}$  years at 90% CL [87]. The hadronic matrix elements for baryon-number-violating operators are evaluated with lattice QCD simulations [88]. The lower bound on the  $X$  boson mass from null results in nucleon decay searches is approaching  $10^{16}$  GeV in SUSY  $SU(5)$  GUTs [89]. In the minimal SUSY  $SU(5)$ ,  $\tau_p / \text{Br}(p \rightarrow K^+ \bar{\nu})$  is smaller than about  $10^{31}$  years if the triplet Higgs mass is  $10^{16}$  GeV and  $m_{\text{SUSY}} = 1$  TeV [90]. The triplet Higgs mass bound from nucleon decay is then in conflict with gauge coupling unification so that this model is considered to be ruled out [68].

Since nucleon decay induced by the triplet Higgs is a severe problem in SUSY GUTs, various proposals for its suppression have been made. First, some accidental symmetry or accidental structure in non-minimal Higgs sectors in  $SU(5)$  or  $SO(10)$  theories may suppress the dimension-five operators [25,26,21,91]. As mentioned above, the triplet Higgs mass term violates symmetries which forbid the dimension-five operators. In other words, the nucleon decay is suppressed if the Higgs triplets in  $\bar{5}_H$  and  $5_H$  do not have a common mass term but, instead, their mass terms involve partners from other  $SU(5)$  multiplets. Second, the SUSY breaking scale may be around  $O(10-100)$  TeV in order to explain the observed Higgs boson mass at the LHC. In this case, nucleon decay is automatically suppressed [72,92,93]. Third, accidental cancellations among diagrams due to a fine-tuned structure of squark and slepton flavor mixing might suppress nucleon decay [94]. Last, we have also implicitly assumed a hierarchical structure for Yukawa matrices in the analysis. It is however possible to fine-tune a hierarchical structure for quarks and leptons which baffles the family structure so that the nucleon decay is suppressed [95]. The upper bound on the proton lifetime from some

of these theories is approximately a factor of 10 above the experimental bounds. Future experiments with larger neutrino detectors, such as JUNO [96], Hyper-Kamiokande [97] and DUNE [98], are planned and will have higher sensitivities to nucleon decay.

Are there ways to avoid the stringent predictions for proton decay discussed above? Orbifold GUTs and string theories, see Sec. 16.4, contain grand unified symmetries realized in higher dimensions. In the process of compactification and GUT symmetry breaking, the triplet Higgs states may be removed (projected out of the massless sector of the theory). In such models, the nucleon decay due to dimension-five operators can be severely suppressed or eliminated completely. However, nucleon decay due to dimension-six operators may be enhanced, since the gauge-bosons mediating proton decay obtain mass at the compactification scale,  $M_c$ , which is typically less than the 4d GUT scale (cf. Sec. 16.5). Alternatively, the same projections which eliminate the triplet Higgs may rearrange the quark and lepton states such that the massless states of one family come from different higher-dimensional GUT multiplets. This can suppress or completely eliminate even dimension-six proton decay. Thus, enhancement or suppression of dimension-six proton decay is model-dependent. In some complete 5-d orbifold GUT models [99,57] the lifetime for the decay  $\tau_p/\text{Br}(p \rightarrow e^+\pi^0)$  can be near the bound of  $1 \times 10^{34}$  years with, however, large model-dependence and/or theoretical uncertainties. In other cases, the modes  $p \rightarrow K^+\bar{\nu}$  and  $p \rightarrow K^0\mu^+$  may be dominant [57]. Thus, interestingly, the observation of nucleon decay may distinguish string or higher-dimensional GUTs from 4d ones.

In orbifold GUTs or string theory, new discrete symmetries consistent with SUSY GUTs can forbid all dimension-three and four baryon- and lepton-number-violating operators. Even the  $\mu$  term and dimension-five baryon- and lepton-number-violating operators can be forbidden to all orders in perturbation theory [32]. The  $\mu$  term and dimension-five baryon- and lepton-number-violating operators may then be generated, albeit sufficiently suppressed, via non-perturbative effects. The simplest example of this is a  $Z_4^R$  symmetry which is the unique discrete  $R$  symmetry consistent with  $SO(10)$  [32]. In this case, nucleon decay is completely dominated by dimension-six operators.

## 16.7. Yukawa Coupling Unification

In the SM, masses and mixings for quarks and leptons come from the Yukawa couplings with the Higgs doublet, but the values of these couplings remain a mystery. GUTs provide at least a partial understanding since each generation is embedded in unified multiplet(s). Specifically, since quarks and leptons are two sides of the same coin, the GUT symmetry relates the Yukawa couplings (and hence the masses) of quarks and leptons.

In  $SU(5)$ , there are two types of independent renormalizable Yukawa interactions given by  $\lambda_{ij}(\mathbf{10}_i \mathbf{10}_j \mathbf{5}_H) + \lambda'_{ij}(\mathbf{10}_i \mathbf{5}_j \mathbf{5}_H)$ . These contain the SM interactions  $\lambda_{ij}(Q_i u_j^c H_u) + \lambda'_{ij}(Q_i d_j^c H_d + e_i^c L_j H_d)$ . Here  $i, j (= 1-3)$  are, as before, family indices. Hence, at the GUT scale we have tree-level relations between Yukawa coupling constants for charged lepton and down quark masses, such as  $\lambda_b = \lambda_\tau$  in which  $\lambda_{b/\tau}$  are the bottom quark /  $\tau$  lepton Yukawa coupling constants [100,101]. In  $SO(10)$ , there is only one type of independent renormalizable Yukawa interaction given by  $\lambda_{ij}(\mathbf{16}_i \mathbf{16}_j \mathbf{10}_H)$ , leading to relations among all Yukawa coupling constants and quark and lepton masses within one generation [102,103] (such as  $\lambda_t = \lambda_b = \lambda_\tau$ , with  $\lambda_t$  the top quark Yukawa coupling constant).

### 16.7.1. The third generation, $b$ - $\tau$ or $t$ - $b$ - $\tau$ unification :

Third generation Yukawa couplings are larger than those of the first two generations. Hence, the fermion mass relations predicted from renormalizable GUT interactions which we introduced above are expected to be more reliable. In order to compare them with data, we have to include the radiative correction to these relations from the RG evolution between GUT and fermion mass scale, from integrating out heavy particles at the GUT scale, and from weak scale thresholds.

Since testing Yukawa coupling unification is only possible in models with successful gauge coupling unification, we here focus on SUSY GUTs. In the MSSM, top and bottom quark and  $\tau$  lepton masses are

related to the Yukawa coupling constants at the scale  $m_Z$  as

$$m_t(m_Z) = \lambda_t(m_Z) v_u(1 + \delta m_t/m_t),$$

$$m_{b/\tau}(m_Z) = \lambda_{b/\tau}(m_Z) v_d(1 + \delta m_{b/\tau}/m_{b/\tau}),$$

where  $\langle H_u^0 \rangle \equiv v_u = \sin \beta v/\sqrt{2}$ ,  $\langle H_d^0 \rangle \equiv v_d = \cos \beta v/\sqrt{2}$ ,  $v_u/v_d \equiv \tan \beta$  and  $v \sim 246$  GeV is fixed by the Fermi constant,  $G_\mu$ . Here,  $\delta m_f/m_f$  ( $f = t, b, \tau$ ) represents the threshold correction due to integrating out SUSY partners. For the bottom quark mass, it is found [104] that the dominant corrections come from the gluino-sbottom and from the Higgsino-stop loops,

$$\left(\frac{\delta m_b}{m_b}\right)_{g_3} \sim \frac{g_3^2}{6\pi^2} \frac{m_{\tilde{g}}\mu}{m_{\text{SUSY}}^2} \tan \beta$$

and

$$\left(\frac{\delta m_b}{m_b}\right)_{\lambda_t} \sim \frac{\lambda_t^2}{16\pi^2} \frac{A_t\mu}{m_{\text{SUSY}}^2} \tan \beta, \quad (16.15)$$

where  $m_{\tilde{g}}$ ,  $\mu$ , and  $A_t$  stand for gluino and Higgsino masses and trilinear stop coupling, respectively. Note that Eq. (16.15) only illustrates the structure of the corrections – non-trivial functional dependences on several soft parameters  $\sim m_{\text{SUSY}}$  have been suppressed. For the full one-loop correction to the bottom quark mass see, for example, Ref. [105].

Note also that the corrections do not go to zero as SUSY particles become much heavier than  $m_Z$ . They may change the bottom quark mass at the 10% level for  $\tan \beta = O(10)$ . The total effect is sensitive to the relative phase between gluino and Higgsino masses since  $A_t \sim -m_{\tilde{g}}$  due to the infrared fixed point nature of the RG equation for  $A_t$  [106] in settings where SUSY breaking terms come from Planck scale dynamics, such as gravity mediation. The  $\tau$  lepton mass also receives a similar correction, though only at the few % level. The top quark mass correction, not being proportional to  $\tan \beta$ , is at most 10% [107].

Including one loop threshold corrections at  $m_Z$  and additional RG running, one finds the top, bottom and  $\tau$  pole masses. In SUSY GUTs,  $b$ - $\tau$  unification has two possible solutions with  $\tan \beta \sim 1$  or  $O(10)$ . The small  $\tan \beta$  solution may be realized in the MSSM if superpartner masses are  $O(10)$  TeV, as suggested by the observed Higgs mass [92]. The large  $\tan \beta$  limit such as  $\tan \beta \sim 40$ –50 overlaps the  $SO(10)$  symmetry relation. When  $\tan \beta$  is large, there are significant threshold correction to down quark mass as mentioned above, and Yukawa unification is only consistent with low-energy data in a restricted region of SUSY parameter space, with important consequences for SUSY searches [107,108]. More recent analyses of Yukawa unification after LHC Run-I are found in Ref. [109].

Gauge coupling unification is also successful in the scenario of split supersymmetry [72], in which squarks and sleptons have mass at a scale  $\tilde{m} \gg m_Z$ , while gauginos and Higgsinos have masses of order the weak scale. Unification of  $b$ - $\tau$  Yukawa couplings requires  $\tan \beta$  to be fine-tuned close to 1 [92]. If by contrast,  $\tan \beta \gtrsim 1.5$ ,  $b$ - $\tau$  Yukawa unification only works for  $\tilde{m} \lesssim 10^4$  GeV. This is because the effective theory between the gaugino mass scale and  $\tilde{m}$  includes only one Higgs doublet, as in the standard model. As a result, the large top quark Yukawa coupling tends to increase the ratio  $\lambda_b/\lambda_\tau$  due to the vertex correction, which is absent in supersymmetric theories, as one runs down in energy below  $\tilde{m}$ . This is opposite to what happens in the MSSM where the large top quark Yukawa coupling lowers the ratio  $\lambda_b/\lambda_\tau$  [101].

### 16.7.2. Beyond leading order: three-family models :

Simple Yukawa unification is not possible for the first two generations. Indeed,  $SU(5)$  implies  $\lambda_s = \lambda_\mu$ ,  $\lambda_d = \lambda_e$  and hence  $\lambda_s/\lambda_d = \lambda_\mu/\lambda_e$ . This is an RG-invariant relation which extrapolates to  $m_s/m_d = m_\mu/m_e$  at the weak scale, in serious disagreement with data ( $m_s/m_d \sim 20$  and  $m_\mu/m_e \sim 200$ ). An elegant solution to this problem was given by Georgi and Jarlskog [110] (for a recent analysis in the SUSY context see [111]).

More generally, we have to recall that in all of the previous discussion of Yukawa couplings, we assumed renormalizable interactions as well



as the minimal matter and Higgs content. Since the GUT scale is close to the Planck scale, higher-dimension operators involving the GUT-breaking Higgs may modify the predictions, especially for lower generations. An example is provided by the operators  $\mathbf{10} \bar{\mathbf{5}} \mathbf{5}_H \mathbf{24}_H$  with  $\mathbf{24}_H$  the GUT-breaking Higgs of  $SU(5)$ . We can fit parameters to the observed fermion masses with these operators, though some fine-tuning is introduced in doing so. The SM Higgs doublet may come in part from higher representations of the GUT group. For example, the  $\mathbf{45}$  of  $SU(5)$  includes an  $SU(2)_L$  doublet with appropriate  $U(1)_Y$  charge [110]. This  $\mathbf{45}$  can, in turn, come from the  $\mathbf{120}$  or  $\mathbf{126}$  of  $SO(10)$  after its breaking to  $SU(5)$  [112]. These fields may also have renormalizable couplings with quarks and leptons. The relations among the Yukawa coupling constants in the SM are modified if the SM Higgs doublet is a linear combination of several such doublets from different  $SU(5)$  multiplets. Finally, the SM fermions may not be embedded in GUT multiplets in the minimal way. Indeed, if all quarks and leptons are embedded in  $\mathbf{16}$ s of  $SO(10)$ , the renormalizable interactions with  $\mathbf{10}_H$  cannot explain the observed CKM mixing angles. This situation improves when extra matter multiplets, such as  $\mathbf{10}$ , are introduced: After  $U(1)_X$ , which distinguishes the  $\bar{\mathbf{5}}$ s coming from the  $\mathbf{16}$  and the  $\mathbf{10}$  of  $SO(10)$ , is broken (e.g. by a VEV of  $\mathbf{16}_H$  or  $\mathbf{126}_H$ ), the r.h. down quarks and l.h. leptons in the SM can be linear combinations of components in  $\mathbf{16}$ s and  $\mathbf{10}$ s. As a result,  $\lambda \neq \lambda'$  in  $SU(5)$  [113].

To construct realistic three-family models, some or all of the above effects can be used. Even so, to achieve significant predictions for fermion masses and mixing angles grand unification alone is not sufficient. Other ingredients, for example additional global family symmetries are needed (in particular, non-abelian symmetries can strongly reduce the number of free parameters). These family symmetries constrain the set of effective higher-dimensional fermion mass operators discussed above. In addition, sequential breaking of the family symmetry can be correlated with the hierarchy of fermion masses. One simple, widely known idea in this context is to ensure that each  $\mathbf{10}_i$  enters Yukawa interactions together with a suppression factor  $\epsilon^{3-i}$  ( $\epsilon$  being a small parameter). This way one automatically generates a stronger hierarchy in up-type quark Yukawas as compared to down-type quark and lepton Yukawas and no hierarchy for neutrinos, which agrees with observations at the  $\mathcal{O}(1)$ -level. Three-family models exist which fit all the data, including neutrino masses and mixing [26,114].

Finally, a particularly ambitious variant of unification is to require that the fermions of all three generations come from a single representation of a large gauge group. A somewhat weaker assumption is that the flavor group (e.g.  $SU(3)$ ) unifies with the SM gauge group in a simple gauge group at some energy scale  $M \geq M_G$ . Early work on such ‘flavor-unified GUTs’, see e.g. [115], has been reviewed in [116,117]. For a selection of more recent papers see [118]. In such settings, Yukawa couplings are generally determined by gauge couplings together with symmetry breaking VEVs. This is reminiscent of heterotic string GUTs, where all couplings come from the 10d gauge coupling. However, while the  $E_8 \rightarrow SU(3) \times E_6$  branching rule  $\mathbf{248} = (\mathbf{8}, \mathbf{1}) + (\mathbf{1}, \mathbf{78}) + (\mathbf{3}, \mathbf{27}) + (\bar{\mathbf{3}}, \mathbf{27})$  looks very suggestive in this context, the way in which most modern heterotic models arrive at three generations is actually more complicated.

**16.7.3. Flavor violation :** Yukawa interactions of GUT-scale particles with quarks and leptons may leave imprints on the flavor violation induced by SUSY breaking parameters [119]. To understand this, focus first on the MSSM with universal Planck-scale boundary conditions (as e.g. in gravity mediation). Working in a basis where up-quark and lepton Yukawas are diagonal, one finds that the large top-quark Yukawa coupling reduces the l.h. squark mass squareds in the third generation radiatively. It turns out that only the l.h. down-type squark mass matrix has sizable off-diagonal terms in the flavor basis after CKM-rotation. However, in GUTs the color-triplet Higgs has flavor violating interactions from the Yukawa coupling  $\lambda_{ij} (\mathbf{10}_i \mathbf{10}_j \mathbf{5}_H)$ , such that flavor-violating r.h. slepton mass terms are radiatively generated in addition [120]. If r.h. neutrinos are introduced as  $SU(5)$  singlets with interactions  $\lambda''_{ij} (\mathbf{1}_i \bar{\mathbf{5}}_j \mathbf{5}_H)$ , the doublet and color-triplet Higgses acquire another type of Yukawa coupling, respectively. They then radiatively generate

flavor-violating l.h. slepton [121] and r.h. down squark masses [122]. These flavor-violating SUSY breaking terms induce new contributions to FCNC processes in quark and lepton sectors, such as  $\mu \rightarrow e\gamma$  and  $K^0-\bar{K}^0$  and  $B^0-\bar{B}^0$  mixing. EDMs are also induced when both l.h. and r.h. squarks/sleptons have flavor-violating mass terms with relative phases, as discussed for  $SO(10)$  in [123] or for  $SU(5)$  with r.h. neutrinos in [124]. Thus, such low-energy observables constrain GUT-scale interactions.

## 16.8. Neutrino Masses

We see from atmospheric and solar neutrino oscillation observations, along with long baseline accelerator and reactor experiments, that neutrinos have finite masses. By adding three ‘sterile’ neutrinos  $\nu_i^c$  with Yukawa couplings  $\lambda_{\nu,ij} (\nu_i^c L_j H_u)$  ( $i, j = 1-3$ ), one easily obtains three massive Dirac neutrinos with mass  $m_\nu = \lambda_\nu v_u$ , analogously to quark and charged lepton masses. However, in order to obtain a  $\tau$  neutrino with mass of order 0.1 eV, one requires the exceedingly small coupling ratio  $\lambda_{\nu\tau}/\lambda_\tau \lesssim 10^{-10}$ . By contrast, the seesaw mechanism *naturally* explains such tiny neutrino masses as follows [1,2,3]: The sterile neutrinos have no SM gauge quantum numbers so that there is no symmetry other than global lepton number which forbids the Majorana mass term  $\frac{1}{2} M_{ij} \nu_i^c \nu_j^c$ . Note also that sterile neutrinos can be identified with the r.h. neutrinos necessarily contained in complete families of  $SO(10)$  or Pati-Salam models. Since the Majorana mass term violates  $U(1)_X$  in  $SO(10)$ , one might expect  $M_{ij} \sim M_G$ . The heavy sterile neutrinos can be integrated out, defining an effective low-energy theory with only three light active Majorana neutrinos with the effective dimension-five operator

$$-\mathcal{L}_{eff} = \frac{1}{2} c_{ij} (L_i H_u) (L_j H_u), \quad (16.16)$$

where  $c = \lambda_\nu^T M^{-1} \lambda_\nu$ . This then leads to a  $3 \times 3$  Majorana neutrino mass matrix  $m = m_\nu^T M^{-1} m_\nu$ .

Atmospheric neutrino oscillations require neutrino masses with  $\Delta m_\nu^2 \sim 2.5 \times 10^{-3} \text{ eV}^2$  with maximal mixing, in the simplest two neutrino scenario. With hierarchical neutrino masses this implies  $m_{\nu\tau} = \sqrt{\Delta m_\nu^2} \sim 0.05 \text{ eV}$ . Next, we can try to relate the neutrino Yukawa coupling to the top quark Yukawa coupling,  $\lambda_{\nu\tau} = \lambda_t$  at the GUT scale, as in  $SO(10)$  or  $SU(4) \times SU(2)_L \times SU(2)_R$  models. This gives  $M \sim 10^{14} \text{ GeV}$ , which is remarkably close to the GUT scale.

Neutrinos pose a special problem for SUSY GUTs. The question is why are the quark mixing angles in the CKM matrix small, while there are two large lepton mixing angles in the PMNS matrix (cf. however the comment at the end of Sec. 16.7). Discussions of neutrino masses and mixing angles can, for example, be found in Refs. [125] and [126]. For SUSY GUT models which fit quark and lepton masses, see Ref. [25]. Finally, for a compilation of the range of SUSY GUT predictions for neutrino mixing, see [127].

The seesaw mechanism implemented by r.h. neutrinos is sometimes called the type-I seesaw model. There are variant models in which the dimension-five operator for neutrino masses is induced in different ways: In the type-II model, an  $SU(2)_L$  triplet Higgs boson  $\Sigma$  is introduced to have couplings  $\Sigma L^2$  and also  $\Sigma H_u^2$  [128]. In the type-III model, an  $SU(2)_L$  triplet of fermions  $\tilde{\Sigma}$  with a Yukawa coupling  $\tilde{\Sigma} L H_u$  is introduced [129]. In these models, the dimension-five operator is induced by integrating out the triplet Higgs boson or fermions. Such models can also be implemented in GUTs by introducing Higgs bosons in the  $\mathbf{15}$  or fermions in the  $\mathbf{24}$  in  $SU(5)$  GUTs or the  $\mathbf{126}$  in  $SO(10)$  GUTs. Notice that the gauge non-singlet fields in the type-II and III models have masses at the intermediate scale. Thus, gauge coupling unification is not automatic if they are implemented in SUSY GUTs.

## 16.9. Selected Topics

### 16.9.1. Magnetic Monopoles :

In the broken phase of a GUT there are typically localized classical solutions carrying magnetic charge under an unbroken  $U(1)$  symmetry [130]. These magnetic monopoles with mass of order  $M_G/\alpha_G$  can be produced during a possible GUT phase transition in the early universe. The flux of magnetic monopoles is experimentally found to be less than  $\sim 10^{-16} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$  [131]. Many more are however predicted, hence the GUT monopole problem. In fact, one of the original motivations for inflation was to solve the monopole problem by exponential expansion after the GUT phase transition [132] and hence dilution of the monopole density. Other possible solutions to the monopole problem include: sweeping them away by domain walls [133],  $U(1)$  electromagnetic symmetry breaking at high temperature [134] or GUT symmetry non-restoration [135]. Parenthetically, it was also shown that GUT monopoles can catalyze nucleon decay [136]. A significantly lower bound on the monopole flux can then be obtained by considering X-ray emission from radio pulsars due to monopole capture and the subsequent nucleon decay catalysis [137].

Note that the present upper bound on the inflationary vacuum energy density is very close to the GUT scale,  $V_{inf}^{1/4} \lesssim M_G$  [138]. This almost guarantees that reheating does not lead to temperatures above  $M_G$  and hence the monopole problem is solved by inflation.

### 16.9.2. Anomaly constraints vs. GUT paradigm :

As emphasized at the very beginning, the fact that the SM fermions of one generation fill out the  $\mathbf{10} + \mathbf{\bar{5}}$  of  $SU(5)$  appears to provide overwhelming evidence for some form of GUT embedding. However, one should be aware that a counterargument can be made which is related to the issue of ‘charge quantization by anomaly cancellation’ (see [139,140] for some early papers and [141] for a more detailed reference list): Imagine we only knew that the low-energy gauge group were  $G_{SM}$  and the matter content included the  $(\mathbf{3}, \mathbf{2})_Y$ , i.e. a ‘quark doublet’ with  $U(1)$ -charge  $Y$ . One can then ask which possibilities exist of adding further matter to ensure the cancellation of all triangle anomalies. It turns out that this problem has only three different, minimal\* solutions [140]. One of those is precisely a single SM generation, with the apparent ‘ $SU(5)$ -ness’ emerging accidentally. Thus, if one randomly picks models from the set of consistent gauge theories, preconditioning on  $G_{SM}$  and  $(\mathbf{3}, \mathbf{2})_Y$ , one may easily end up with ‘ $\mathbf{10} + \mathbf{\bar{5}}$ ’ of an  $SU(5)$  that is in no way dynamically present. This is precisely what happens in the context of non-GUT string model building [142].

### 16.9.3. GUT Baryogenesis and Leptogenesis :

Baryon-number-violating operators in  $SU(5)$  or  $SO(10)$  preserve the global symmetry  $(B - L)$ . Hence the value of the cosmological  $(B - L)$  density is an initial condition of the theory and is typically assumed to be zero. On the other hand, anomalies of the electroweak symmetry violate  $(B + L)$  while also preserving  $(B - L)$ . Hence thermal fluctuations in the early universe, via so-called sphaleron processes, can drive  $(B + L)$  to zero, washing out any net baryon number generated in the early universe at GUT temperatures. In particular, this affects the old idea of GUT baryogenesis [143,144], where a  $(B + L)$  asymmetry is generated by the out-of-equilibrium decay of the color-triplet Higgs. A possible way out [145] uses lepton-number violating interaction of neutrinos to create a  $(B - L)$  asymmetry from the  $(B + L)$  symmetry, before sphaleron processes become sufficiently fast at  $T < 10^{12} \text{ GeV}$ . This  $(B - L)$  asymmetry can then survive the subsequent sphaleron dominated phase. Note that this does not work in the minimal SUSY GUT setting, with the triplet Higgs above the GUT scale. The reason is that a correspondingly high reheating temperature would be required, leading to monopole overproduction.

However, the most widely accepted simple way out of the dilemma is to directly generate a net  $(B - L)$  asymmetry dynamically in the early universe, also using r.h. neutrinos. Indeed, we have seen that

neutrino oscillations suggest a new scale of physics of order  $10^{14} \text{ GeV}$ . This scale is associated with heavy Majorana neutrinos in the seesaw mechanism. If in the early universe, the decay of the heavy neutrinos is out of equilibrium and violates both lepton number and CP, then a net lepton number may be generated. This lepton number will then be partially converted into baryon number via electroweak processes [146]. This mechanism is called leptogenesis.

If the three heavy Majorana neutrino masses are hierarchical, the net lepton number is produced by decay of the lightest one, and it is proportional to the CP asymmetry in the decay,  $\epsilon_1$ . The CP asymmetry is bounded from above, and the lightest neutrino mass is required to be larger than  $10^9 \text{ GeV}$  in order to explain the observed baryon asymmetry [147]. This implies that the reheating temperature after inflation should be larger than  $10^9 \text{ GeV}$  so that the heavy neutrinos are thermally produced. In supersymmetric models, there is a tension between leptogenesis and Big Bang Nucleosynthesis (BBN) if gravitinos decay in the BBN era. The gravitino problem gives a constraint on the reheating temperature  $\lesssim 10^{6-10} \text{ GeV}$  though the precise value depends on the SUSY breaking parameters [148]. Recent reviews of leptogenesis can be found in Ref. [149].

## 16.10. Conclusion

Most conservatively, grand unification means that (some of) the SM gauge interactions of  $U(1)_Y$ ,  $SU(2)_L$  and  $SU(3)_C$  become part of a larger, unifying gauge symmetry at a high energy scale. In most models, especially in the simplest and most appealing variants of  $SU(5)$  and  $SO(10)$  unification, the statement is much stronger: One expects the three gauge couplings to unify (up to small threshold corrections) at a unique scale,  $M_G$ , and the proton to be unstable due to exchange of gauge bosons of the larger symmetry group. Supersymmetric grand unified theories provide, by far, the most predictive and economical framework allowing for perturbative unification. For a selection of reviews, with many more details than could be discussed in the present article, see [116,150].

Thus, the three classical pillars of GUTs are gauge coupling unification at  $M_G \sim 2 \times 10^{16} \text{ GeV}$ , low-energy supersymmetry (with a large SUSY desert), and nucleon decay. The first of these may be viewed as predicting the value of the strong coupling – a prediction which has already been verified (see Fig. 16.1). While this remains true even if SUSY partner masses are somewhat above the weak scale, the possible complete absence of SUSY in the LHC energy range is nevertheless problematic for the GUT paradigm: If the independent, gauge-hierarchy-based motivation for SUSY is completely abandoned, the SUSY scale and hence  $\alpha_3$  become simply free parameters and the first two pillars crumble. It is the more important to keep pushing bounds on proton decay which, although again not completely universal in all GUT constructions, is arguably a more generic part of the GUT paradigm than low-energy SUSY.

Whether or not Yukawa couplings unify is more model dependent. However, irrespective of possible (partial) Yukawa unification, there certainly exists a very interesting and potentially fruitful interplay between flavor model building and grand unification. Especially in the neutrino sector this is strongly influenced by the developing experimental situation.

It is probably fair to say that, due to limitations of the 4d approach, including especially remaining ambiguities (free parameters or ad hoc assumptions) in models of flavor and GUT breaking, the string theoretic approach has become more important in GUT model building. In this framework, challenges include learning how to deal with the many vacua of the ‘landscape’ as well as, for each vacuum, developing the tools for reliably calculating detailed, phenomenological observables. Finally, due to limitations of space, the present article has barely touched on the interesting cosmological implications of GUTs. They may become more important in the future, especially in the case that a high inflationary energy scale is established observationally.

## References:

1. P. Minkowski, Phys. Lett. **B67**, 421 (1977).
2. T. Yanagida, in *Proceedings of the Workshop on the Unified Theory and the Baryon Number of the Universe*, eds. O. Sawada, A. Sugamoto, KEK report No. 79-18, Tsukuba, Japan, 1979;

\* Adding extra vector-like sets of fields, e.g. two fermions which only transform under  $U(1)$  and have charges  $Y$  and  $-Y$ , is considered to violate minimality.

- S. Glashow, Quarks and leptons, published in *Proceedings of the Cargèse Lectures*, M. Levy (ed.), Plenum Press, New York, (1980);  
M. Gell-Mann, P. Ramond and R. Slansky, in *Supergravity*, ed. P. van Nieuwenhuizen *et al.*, North-Holland, Amsterdam, (1979), p. 315 [arXiv:1306.4669].
3. F. Wilczek, eConf C **790823**, 437 (1979);  
E. Witten, Phys. Lett. **B91**, 81 (1980);  
R.N. Mohapatra and G. Senjanovic, Phys. Rev. Lett. **44**, 912 (1980).
4. U. Amaldi, W. de Boer and H. Furstenau, Phys. Lett. **B260**, 447 (1991);  
J.R. Ellis, S. Kelley and D.V. Nanopoulos, Phys. Lett. **B260**, 131 (1991);  
P. Langacker and M. Luo, Phys. Rev. **D44**, 817 (1991);  
C. Giunti, C.W. Kim and U.W. Lee, Mod. Phys. Lett. **A6**, 1745 (1991);  
P. Langacker and N. Polonsky, Phys. Rev. **D47**, 4028 (1993), [hep-ph/9210235];  
M. Carena, S. Pokorski and C.E.M. Wagner, Nucl. Phys. **B406**, 59 (1993) [hep-ph/9303202] see also the review by S. Dimopoulos, S.A. Raby and F. Wilczek, Physics Today, p. 25 October (1991).
5. Y. Okada, M. Yamaguchi and T. Yanagida, Prog. Theor. Phys. **85**, 1 (1991);  
J.R. Ellis, G. Ridolfi and F. Zwirner, Phys. Lett. **B257**, 83 (1991);  
H.E. Haber and R. Hempfling, Phys. Rev. Lett. **66**, 1815 (1991).
6. M.S. Carena *et al.*, Phys. Lett. **B355**, 209 (1995) [hep-ph/9504316];  
G. Degrandi *et al.*, Eur. Phys. J. **C28**, 133 (2003) [arXiv:hep-ph/0212020];  
P. Kant *et al.*, JHEP **1008**, 104 (2010) [arXiv:1005.5709].
7. M.J.G. Veltman, Acta Phys. Pol. **B12**, 437 (1981);  
L. Maiani, Gif-sur-Yvette Summer School on Particle Physics, 11th, Gif-sur-Yvette, France, 1979 (Inst. Nat. Phys. Nucl. Phys. Particules, Paris, 1979);  
E. Witten, Nucl. Phys. **B188**, 513 (1981).
8. J.L. Feng *et al.*, Phys. Rev. Lett. **111**, 131802 (2013) [arXiv:1306.2318];  
S. Heinemeyer, Lecture at 42nd ITEP winter school, Moscow, arXiv:1405.3781;  
P. Draper and H. Rzehak, arXiv:1601.01890.
9. J. Pati and A. Salam, Phys. Rev. **D8**, 1240 (1973);  
For more discussion on the standard charge assignments in this formalism, see A. Davidson, Phys. Rev. **D20**, 776 (1979) and R.N. Mohapatra and R.E. Marshak, Phys. Lett. **B91**, 222 (1980).
10. H. Georgi and S.L. Glashow, Phys. Rev. Lett. **32**, 438 (1974).
11. J.R. Ellis, M.K. Gaillard and D.V. Nanopoulos, Phys. Lett. **B80**, 360 (1979) [Phys. Lett. **B82**, 464 (1979)];  
E. Golowich, Phys. Rev. **D24**, 2899 (1981).
12. S. Dimopoulos, S. Raby, and F. Wilczek, Phys. Rev. **D24**, 1681 (1981);  
S. Dimopoulos and H. Georgi, Nucl. Phys. **B193**, 150 (1981);  
L. Ibanez and G.G. Ross, Phys. Lett. **B105**, 439 (1981);  
N. Sakai, Z. Phys. **C11**, 153 (1981);  
M.B. Einhorn and D.R.T. Jones, Nucl. Phys. **B196**, 475 (1982);  
W.J. Marciano and G. Senjanovic, Phys. Rev. **D25**, 3092 (1982).
13. S. Weinberg, Phys. Rev. **D26**, 287 (1982);  
N. Sakai and T. Yanagida, Nucl. Phys. **B197**, 533 (1982).
14. H. Georgi, Particles and Fields, *Proceedings of the APS Div. of Particles and Fields*, ed. C. Carlson, p. 575 (1975);  
H. Fritzsch and P. Minkowski, Ann. Phys. **93**, 193 (1975).
15. R. Slansky, Phys. Reports **79**, 1 (1981);  
H. Georgi, Front. Phys. **54**, 1 (1982);  
R. Feger and T.W. Kephart, Comp. Phys. Comm. **192**, 166 (2015) [arXiv:1206.6379 [math-ph]];  
N. Yamatsu, preprint OU-HET 888 [arXiv:1512.05559].
16. S.M. Barr, Phys. Lett. **B112**, 219 (1982);  
J.P. Derendinger, J.E. Kim and D.V. Nanopoulos, Phys. Lett. **B139**, 170 (1984);  
I. Antoniadis *et al.*, Phys. Lett. **B194**, 231 (1987) and Phys. Lett. **B231**, 65 (1989).
17. F. Gursey, P. Ramond and P. Sikivie, Phys. Lett. **B60**, 177 (1976).
18. A. de Rujula, H. Georgi and S. L. Glashow, *5th Workshop on Grand Unification*, ed. K. Kang, H. Fried and P. Frampton, World Scientific, Singapore (1984), p. 88;  
See also earlier paper by Y. Achiman and B. Stech, p. 303, "New Phenomena in Lepton-Hadron Physics," ed. D.E.C. Fries and J. Wess, Plenum, NY (1979).
19. B.R. Greene *et al.*, Nucl. Phys. **B278**, 667 (1986) and Nucl. Phys. **B292**, 606 (1987);  
B.R. Greene, C.A. Lutken and G.G. Ross, Nucl. Phys. **B325**, 101 (1989);  
J.E. Kim, Phys. Lett. **B591**, 119 (2004) [hep-ph/0403196].
20. E. Witten, Phys. Lett. **B105**, 267 (1981).
21. A. Masiero *et al.*, Phys. Lett. **B115**, 380 (1982);  
B. Grinstein, Nucl. Phys. **B206**, 387 (1982).
22. S. Dimopoulos and F. Wilczek, *Proceedings Erice Summer School*, ed. A. Zichichi (1981);  
M. Srednicki, Nucl. Phys. **B202**, 327 (1982).
23. K. Inoue, A. Kakuto and H. Takano, Prog. Theor. Phys. **75**, 664 (1986).
24. Y. Yamada, Z. Phys. **C60**, 83 (1993);  
J. Hisano *et al.*, Phys. Lett. **B342**, 138 (1995) [hep-ph/9406417];  
G. Altarelli, F. Feruglio and I. Masina, JHEP **0011**, 040 (2000) [hep-ph/0007254].
25. K.S. Babu and S.M. Barr, Phys. Rev. **D48**, 5354 (1993) [hep-ph/9306242];  
K.S. Babu and S.M. Barr, Phys. Rev. **D50**, 3529 (1994) [hep-ph/9402291];  
K.S. Babu, J.C. Pati and Z. Tavartkiladze, JHEP **1006**, 084 (2010) [arXiv:1003.2625].
26. K.S. Babu and R.N. Mohapatra, Phys. Rev. Lett. **74**, 2418 (1995) [hep-ph/9410326];  
V. Lucas and S. Raby, Phys. Rev. **D54**, 2261 (1996) [hep-ph/9601303];  
T. Blažek *et al.*, Phys. Rev. **D56**, 6919 (1997) [hep-ph/9611217];  
S.M. Barr and S. Raby, Phys. Rev. Lett. **79**, 4748 (1997) [hep-ph/9705366];  
K.S. Babu, J.C. Pati and F. Wilczek, Nucl. Phys. **B566**, 33 (2000) [hep-ph/9812538];  
R. Dermisek, A. Mafi and S. Raby, Phys. Rev. **D63**, 035001 (2001) [hep-ph/0007213].
27. R. Barbieri, G.R. Dvali and A. Strumia, Nucl. Phys. **B391**, 487 (1993);  
Z. Berezhiani, C. Csaki and L. Randall, Nucl. Phys. **B444**, 61 (1995) [hep-ph/9501336];  
Q. Shafi and Z. Tavartkiladze, Phys. Lett. **B522**, 102 (2001) [hep-ph/0105140].
28. G.F. Giudice and A. Masiero, Phys. Lett. **B206**, 480 (1988);  
J.E. Kim and H.P. Nilles, Mod. Phys. Lett. **A9**, 3575 (1994) [hep-ph/9406296].
29. L. Randall and C. Csaki, in *Palaiseau 1995, SUSY 95* 99-109 [hep-ph/9508208].
30. E. Witten, [hep-ph/0201018] M. Dine, Y. Nir and Y. Shadmi, Phys. Rev. **D66**, 115001 (2002) [hep-ph/0206268].
31. A. Hebecker, J. March-Russell and R. Ziegler, JHEP **0908**, 064 (2009) [arXiv:0801.4101];  
F. Brümmer *et al.*, JHEP **0908**, 011 (2009) [arXiv:0906.2957];  
F. Brümmer *et al.*, JHEP **1004**, 006 (2010) [arXiv:1003.0084].
32. H.M. Lee *et al.*, Phys. Lett. **B694**, 491 (2011) [arXiv:1009.0905];  
R. Kappl *et al.*, Nucl. Phys. **B847**, 325 (2011) [arXiv:1012.4574];  
H.M. Lee *et al.*, Nucl. Phys. **B850**, 1 (2011) [arXiv:1102.3595].
33. Y. Hosotani, Phys. Lett. **B126**, 309 (1983).
34. L.J. Dixon *et al.*, Nucl. Phys. **B261**, 678 (1985) and Nucl. Phys. **B274**, 285 (1986);

- L.E. Ibanez, H.P. Nilles and F. Quevedo, Phys. Lett. **B187**, 25 (1987);  
 L.E. Ibanez *et al.*, Phys. Lett. **B191**, 282 (1987).
35. P. Candelas *et al.*, Nucl. Phys. **B258**, 46 (1985).
  36. Y. Kawamura, Prog. Theor. Phys. **103**, 613 (2000) [hep-ph/9902423] and Prog. Theor. Phys. **105**, 999 (2001) [hep-ph/0012125];  
 G. Altarelli and F. Feruglio, Phys. Lett. **B511**, 257 (2001) [hep-ph/0102301];  
 L.J. Hall and Y. Nomura, Phys. Rev. **D64**, 055003 (2001) [hep-ph/0103125];  
 A. Hebecker and J. March-Russell, Nucl. Phys. **B613**, 3 (2001) [hep-ph/0106166];  
 T. Asaka, W. Buchmuller and L. Covi, Phys. Lett. **B523**, 199 (2001) [hep-ph/0108021];  
 L.J. Hall *et al.*, Phys. Rev. **D65**, 035008 (2002) [hep-ph/0108071];  
 R. Dermisek and A. Mafi, Phys. Rev. **D65**, 055002 (2002) [hep-ph/0108139];  
 H.D. Kim and S. Raby, JHEP **0301**, 056 (2003) [hep-ph/0212348].
  37. L.E. Ibanez and A.M. Uranga, "String theory and particle physics: An introduction to string phenomenology," Cambridge University Press 2012 K.S. Choi and J.E. Kim, Lect. Notes Phys. **696**, 1 (2006) R. Blumenhagen *et al.*, Phys. Reports **445**, 1 (2007) [hep-th/0610327].
  38. V. Braun *et al.*, JHEP **0605**, 043 (2006) [hep-th/0512177];  
 V. Bouchard and R. Donagi, Phys. Lett. **B633**, 783 (2006) [hep-th/0512149];  
 L.B. Anderson *et al.*, Phys. Rev. **D84**, 106005 (2011) [arXiv:1106.4804].
  39. L.J. Dixon, V. Kaplunovsky and J. Louis, Nucl. Phys. **B355**, 649 (1991).
  40. G. Aldazabal *et al.*, Nucl. Phys. **B452**, 3 (1995) [hep-th/9410206];  
 Z. Kakushadze *et al.*, Int. J. Mod. Phys. **A13**, 2551 (1998) [hep-th/9710149].
  41. E. Witten, Nucl. Phys. **B258**, 75 (1985).
  42. S. Gukov *et al.*, Phys. Rev. **D69**, 086008 (2004) [hep-th/0310159];  
 G. Curio, A. Krause and D. Lust, Fortsch. Phys. **54**, 225 (2006) [hep-th/0502168];  
 L.B. Anderson *et al.*, Phys. Rev. **D83**, 106011 (2011) [arXiv:1102.0011].
  43. R. Friedman, J. Morgan and E. Witten, Commun. Math. Phys. **187**, 679 (1997) [hep-th/9701162].
  44. R. Blumenhagen, G. Honecker and T. Weigand, JHEP **0508**, 009 (2005) [hep-th/0507041].
  45. T. Kobayashi, S. Raby and R.J. Zhang, Phys. Lett. **B593**, 262 (2004) [hep-ph/0403065];  
 S. Forste *et al.*, Phys. Rev. **D70**, 106008 (2004) [hep-th/0406208];  
 T. Kobayashi, S. Raby and R.J. Zhang, Nucl. Phys. **B704**, 3 (2005) [hep-ph/0409098];  
 W. Buchmuller *et al.*, Nucl. Phys. **B712**, 139 (2005) [hep-ph/0412318];  
 W. Buchmuller *et al.*, Phys. Rev. Lett. **96**, 121602 (2006) [hep-ph/0511035], and Nucl. Phys. **B785**, 149 (2007) [hep-th/0606187];  
 O. Lebedev, *et al.*, Phys. Lett. **B645**, 88 (2007) [hep-th/0611095];  
 J.E. Kim, J.H. Kim and B. Kyae, JHEP **0706**, 034 (2007) [hep-ph/0702278];  
 O. Lebedev, *et al.*, Phys. Rev. **D77**, 046013 (2008) [arXiv:0708.2691].
  46. S.B. Giddings, S. Kachru and J. Polchinski, Phys. Rev. **D66**, 106006 (2002) [hep-th/0105097];  
 S. Kachru *et al.*, Phys. Rev. **D68**, 046005 (2003) [hep-th/0301240].
  47. R. Blumenhagen *et al.*, Nucl. Phys. **B815**, 1 (2009) [arXiv:0811.2936].
  48. R. Blumenhagen *et al.*, Nucl. Phys. **B616**, 3 (2001) [hep-th/0107138].
  49. R. Donagi and M. Wijnholt, Adv. Theor. Math. Phys. **15**, 1237 (2011) [arXiv:0802.2969].
  50. C. Beasley, J.J. Heckman and C. Vafa, JHEP **0901**, 058 (2009) [arXiv:0802.3391], and JHEP **0901**, 059 (2009) [arXiv:0806.0102].
  51. T. Weigand, Class. Quantum Grav. **27**, 214004 (2010) [arXiv:1009.3497] J.J. Heckman, Ann. Rev. Nucl. and Part. Sci. **160**, 237 (2010) [arXiv:1001.0577] M. Cvetič, I. Garcia-Etxebarria and J. Halverson, JHEP **1101**, 073 (2011) [arXiv:1003.5337];  
 A. Maharana and E. Palti, Int. J. Mod. Phys. **A28**, 1330005 (2013) [arXiv:1212.0555];  
 S. Krippendorff, S. Schafer-Nameki and J.M. Wong, JHEP **1511**, 008 (2015) [arXiv:1507.05961].
  52. R. Donagi and M. Wijnholt, Adv. Theor. Math. Phys. **15**, 1523 (2011) [arXiv:0808.2223];  
 R. Blumenhagen, Phys. Rev. Lett. **102**, 071601 (2009) [arXiv:0812.0248];  
 C. Mayrhofer, E. Palti and T. Weigand, JHEP **1309**, 082 (2013) [arXiv:1303.3589].
  53. E. Witten, Nucl. Phys. **B471**, 135 (1996) [hep-th/9602070].
  54. A. Hebecker and M. Trapletti, Nucl. Phys. **B713**, 173 (2005) [hep-th/0411131].
  55. H. Georgi, H. R. Quinn and S. Weinberg, Phys. Rev. Lett. **33**, 451 (1974) See also the definition of effective field theories by S. Weinberg, Phys. Lett. **B91**, 51 (1980).
  56. K.A. Olive *et al.*, [Particle Data Group], Chin. Phys. C **38**, 090001 (2014).
  57. M.L. Alciati *et al.*, JHEP **0503**, 054 (2005) [hep-ph/0501086].
  58. See talks on proposed and running nucleon decay experiments, and theoretical talks by P. Langacker, p. 131, and W.J. Marciano and A. Sirlin, p. 151, in *The Second Workshop on Grand Unification*, eds. J.P. Leveille *et al.*, Birkhäuser, Boston (1981).
  59. W.J. Marciano, p. 190, *Eighth Workshop on Grand Unification*, ed. K. Wali, World Scientific Publishing Co., Singapore (1987).
  60. M.S. Carena *et al.*, in Ref. [4].
  61. B.C. Allanach, Comp. Phys. Comm. **143**, 305 (2002) [hep-ph/0104145].
  62. A. Djouadi, J. L. Kneur and G. Moultaka, Comp. Phys. Comm. **176**, 426 (2007) [hep-ph/0211331];  
 W. Porod and F. Staub, Comp. Phys. Comm. **183**, 2458 (2012) [arXiv:1104.1573].
  63. W. Siegel, Phys. Lett. **B94**, 37 (1980).
  64. I. Antoniadis, C. Kounnas, and R. Lacaze, Nucl. Phys. **B221**, 377 (1983).
  65. M.A. Shifman, Int. J. Mod. Phys. **A11**, 5761 (1996) [arXiv:hep-ph/9606281];  
 N. Arkani-Hamed and H. Murayama, JHEP **0006**, 030 (2000) [arXiv:hep-th/9707133];  
 I. Jack, D.R.T. Jones and A. Pickering, Phys. Lett. **B435**, 61 (1998) [arXiv:hep-ph/9805482].
  66. M.B. Einhorn and D.R.T. Jones in Ref. [12], I. Antoniadis, C. Kounnas, and K. Tamvakis, Phys. Lett. **B119**, 377 (1982).
  67. J. Hisano, H. Murayama and T. Yanagida, Phys. Rev. Lett. **69**, 1014 (1992) and Nucl. Phys. **B402**, 46 (1993) [hep-ph/9207279].
  68. H. Murayama and A. Pierce, Phys. Rev. **D65**, 055009 (2002) [hep-ph/0108104].
  69. G. Anderson *et al.*, eConf **C960625**, SUP107 (1996) [hep-ph/9609457].
  70. S. Raby, M. Ratz and K. Schmidt-Hoberg, Phys. Lett. **B687**, 342 (2010) [arXiv:0911.4249].
  71. J. Hisano, T. Kuwahara and N. Nagata, Phys. Lett. **B723**, 324 (2013) [arXiv:1304.0343].
  72. N. Arkani-Hamed, A. Delgado and G.F. Giudice, Nucl. Phys. **B741**, 108 (2006) [hep-ph/0601041].
  73. M. Ibe, T. Moroi and T.T. Yanagida, Phys. Lett. **B644**, 355 (2007) [hep-ph/0610277].

74. G.F. Giudice *et al.*, JHEP **9812**, 027 (1998) [hep-ph/9810442]; L. Randall and R. Sundrum, Nucl. Phys. **B557**, 79 (1999) [hep-th/9810155].
75. R.N. Mohapatra and M.K. Parida, Phys. Rev. **D47**, 264 (1993) [hep-ph/9204234].
76. D.G. Lee *et al.*, Phys. Rev. **D51**, 229 (1995) [hep-ph/9404238].
77. L.E. Ibanez *et al.*, JHEP **1207**, 195 (2012) [arXiv:1206.2655].
78. K.R. Dienes, E. Dudas and T. Gherghetta, Phys. Rev. Lett. **91**, 061601 (2003) [hep-th/0210294].
79. K. Agashe, R. Contino and R. Sundrum, Phys. Rev. Lett. **95**, 171804 (2005) [hep-ph/0502222].
80. S. Weinberg, Phys. Rev. Lett. **43**, 1566 (1979); F. Wilczek and A. Zee, Phys. Rev. Lett. **43**, 1571 (1979).
81. R. Barbier *et al.*, Phys. Rev. **D70**, 1 (2005) [hep-ph/0406039].
82. G. Farrar and P. Fayet, Phys. Lett. **B76**, 575 (1978).
83. S. Dimopoulos, S. Raby and F. Wilczek, Phys. Lett. **B112**, 133 (1982); J. Ellis, D.V. Nanopoulos and S. Rudaz, Nucl. Phys. **B202**, 43 (1982).
84. For a recent discussion, see C.S. Aulakh *et al.*, Nucl. Phys. **B597**, 89 (2001) [hep-ph/0004031].
85. R. Kallosh *et al.*, Phys. Rev. **D52**, 912 (1995) [hep-th/9502069].
86. L.E. Ibanez and G.G. Ross, Nucl. Phys. **B368**, 3 (1992); M. Berasaluce-Gonzalez *et al.*, JHEP **1112**, 113 (2011) [arXiv:1106.4169 [hep-th]].
87. M. Ikeda [Super-Kamiokande Collab.], NNN2015 (<https://www.bnl.gov/nnn2015/>).
88. Y. Aoki *et al.*, Phys. Rev. **D89**, 014505 (2014) [arXiv:1304.7424].
89. For recent analysis, see J. Hisano, T. Kuwahara and Y. Omura, Nucl. Phys. **B898**, 1 (2015) [arXiv:1503.08561].
90. T. Goto and T. Nihei, Phys. Rev. **D59**, 115009 (1999) [hep-ph/9808255].
91. J.L. Chkareuli and I.G. Gogoladze, Phys. Rev. **D58**, 055011 (1998) [hep-ph/9803335].
92. G.F. Giudice and A. Romanino, Nucl. Phys. **B699**, 65 (2004) [hep-ph/0406088], [Erratum: *ibid.*, Nucl. Phys. **B706**, 65 (2005)].
93. J. Hisano *et al.*, JHEP **1307**, 038 (2013) [arXiv:1304.3651].
94. B. Bajc, P. Fileviez Perez and G. Senjanovic, Phys. Rev. **D66**, 075005 (2002) [hep-ph/0204311].
95. K.S. Choi, Phys. Lett. **B668**, 392 (2008) [arXiv:0807.2766].
96. Z. Djuricic *et al.* [JUNO Collab.], “JUNO Conceptual Design Report,” [arXiv:1508.07166 [physics.ins-det]].
97. K. Abe *et al.*, “Letter of Intent: The Hyper-Kamiokande Experiment — Detector Design and Physics Potential —,” [arXiv:1109.3262].
98. M. Bishai *et al.*, “Long-Baseline Neutrino Facility (LBNF) and Deep Underground Neutrino Experiment (DUNE) — Conceptual Design Report —.”
99. L.J. Hall and Y. Nomura, Phys. Rev. **D66**, 075004 (2002) [hep-ph/0205067]; H.D. Kim, S. Raby and L. Schradin, JHEP **0505**, 036 (2005) [hep-ph/0411328].
100. M.S. Chanowitz, J.R. Ellis and M.K. Gaillard, Nucl. Phys. **B128**, 506 (1977) A.J. Buras *et al.*, Nucl. Phys. **B135**, 66 (1978).
101. K. Inoue *et al.*, Prog. Theor. Phys. **67**, 1889 (1982); L.E. Ibanez and C. Lopez, Phys. Lett. **B126**, 54 (1983) and Nucl. Phys. **B233**, 511 (1984).
102. H. Georgi and D.V. Nanopoulos, Nucl. Phys. **B159**, 16 (1979); J. Harvey, P. Ramond and D.B. Reiss, Phys. Lett. **B92**, 309 (1980) and Nucl. Phys. **B199**, 223 (1982).
103. T. Banks, Nucl. Phys. **B303**, 172 (1988); M. Olechowski and S. Pokorski, Phys. Lett. **B214**, 393 (1988); S. Pokorski, Nucl. Phys. (Proc. Supp.) **B13**, 606 (1990); B. Ananthanarayan, G. Lazarides and Q. Shafi, Phys. Rev. **D44**, 1613 (1991); Q. Shafi and B. Ananthanarayan, ICTP Summer School lectures (1991); S. Dimopoulos, L.J. Hall and S. Raby, Phys. Rev. Lett. **68**, 1984 (1992) and Phys. Rev. **D45**, 4192 (1992); G. Anderson *et al.*, Phys. Rev. **D47**, 3702 (1993) [hep-ph/9209250]; B. Ananthanarayan, G. Lazarides and Q. Shafi, Phys. Lett. **B300**, 245 (1993); G. Anderson *et al.*, Phys. Rev. **D49**, 3660 (1994) [hep-ph/9308333]; B. Ananthanarayan, Q. Shafi and X.M. Wang, Phys. Rev. **D50**, 5980 (1994) [hep-ph/9311225].
104. L.J. Hall, R. Rattazzi and U. Sarid, Phys. Rev. **D50**, 7048 (1994) [hep-ph/9306309, hep-ph/9306309]; M. Carena *et al.*, Nucl. Phys. **B426**, 269 (1994) [hep-ph/9402253].
105. A. Anandakrishnan, B.C. Bryant and S. Raby, JHEP **1505**, 088 (2015) [arXiv:1411.7035].
106. M. Lanzagorta and G.G. Ross, Phys. Lett. **B364**, 163 (1995) [hep-ph/9507366].
107. K. Tobe and J.D. Wells, Nucl. Phys. **B663**, 123 (2003) [hep-ph/0301015].
108. T. Blazek, R. Dermisek and S. Raby, Phys. Rev. Lett. **88**, 111804 (2002) and Phys. Rev. **D65**, 115004 (2002) [hep-ph/0107097]; D. Auto, *et al.*, JHEP **0306**, 023 (2003) [hep-ph/0302155]; R. Dermisek *et al.*, JHEP **0304**, 037 (2003) and JHEP **0509**, 029 (2005) [hep-ph/0304101].
109. A. Anandakrishnan, S. Raby and A. Wingerter, Phys. Rev. **D87**, 055005 (2013) [arXiv:1212.0542]; M. Adeel Ajaib *et al.*, JHEP **1307**, 139 (2013) [arXiv:1303.6964]; Z. Poh and S. Raby, Phys. Rev. **D92**, 015017 (2015) [arXiv:1505.00264]; M. Badziak, M. Olechowski and S. Pokorski, JHEP **1310**, 088 (2013) [arXiv:1307.7999].
110. H. Georgi and C. Jarlskog, Phys. Lett. **B86**, 297 (1979).
111. S. Antusch and M. Spinrath, Phys. Rev. **D79**, 095004 (2009) [arXiv:0902.4644].
112. G. Lazarides, Q. Shafi and C. Wetterich, Nucl. Phys. **B181**, 287 (1981); T.E. Clark, T.K. Kuo and N. Nakagawa, Phys. Lett. **B115**, 26 (1982); K.S. Babu and R.N. Mohapatra, Phys. Rev. Lett. **70**, 2845 (1993) [hep-ph/9209215].
113. R. Barbieri and D.V. Nanopoulos, Phys. Lett. **B91**, 369 (1980).
114. R. Barbieri and A. Strumia, Nucl. Phys. **B493**, 3 (1997) [hep-ph/9704402]; T. Blazek, S. Raby and K. Tobe, Phys. Rev. **D60**, 113001 (1999) and Phys. Rev. **D62**, 055001 (2000) [hep-ph/9903340]; Q. Shafi and Z. Tavartkiladze, Phys. Lett. **B487**, 145 (2000) [hep-ph/0002150]; C.H. Albright and S.M. Barr, Phys. Rev. Lett. **85**, 244 (2000) [hep-ph/0002155]; Z. Berezhiani and A. Rossi, Nucl. Phys. **B594**, 113 (2001) [hep-ph/0003084]; C.H. Albright and S.M. Barr, Phys. Rev. **D64**, 073010 (2001) [hep-ph/0104294]; M.-C. Chen, K.T. Mahanthappa, Int. J. Mod. Phys. **A18**, 5819 (2003) [hep-ph/0305088]; R. Dermisek and S. Raby, Phys. Lett. **B622**, 327 (2005) [hep-ph/0507045].
115. H. Georgi, Nucl. Phys. **B156**, 126 (1979); P.H. Frampton, Phys. Lett. **B88**, 299 (1979); P. Frampton and S. Nandi, Phys. Rev. Lett. **43**, 1460 (1979); J.E. Kim, Phys. Rev. Lett. **45**, 1916 (1980), Phys. Rev. **D23**, 2706 (1981) and Phys. Rev. **D26**, 674 (1982); R. Barbieri and D.V. Nanopoulos, Phys. Lett. **B91**, 369 (1980); Y. Fujimoto, Phys. Rev. **D26**, 3183 (1982).
116. P. Langacker, Phys. Reports **72**, 185 (1981).
117. H. Georgi, Conf. Proc. C **820726**, 705 (1982).
118. S.M. Barr, Phys. Rev. **D78**, 055008 (2008) [arXiv:0805.4808 [hep-ph]]; Y. Goto, Y. Kawamura and T. Miura, Phys. Rev. **D88**, 055016 (2013) [arXiv:1307.2631]; J.E. Kim, JHEP **1506**, 114 (2015) [arXiv:1503.03104 [hep-ph]];

- C.H. Albright, R.P. Feger and T.W. Kephart, arXiv:1601.07523 [hep-ph].
119. L.J. Hall, V.A. Kostelecky and S. Raby, Nucl. Phys. **B267**, 415 (1986).
  120. R. Barbieri and L.J. Hall, Phys. Lett. **B338**, 212 (1994) [hep-ph/9408406];  
R. Barbieri, L.J. Hall and A. Strumia, Nucl. Phys. **B445**, 219 (1995) [hep-ph/9501334].
  121. F. Borzumati and A. Masiero, Phys. Rev. Lett. **57**, 961 (1986);  
J. Hisano *et al.*, Phys. Lett. **B357**, 579 (1995) [hep-ph/9501407];  
J. Hisano *et al.*, Phys. Rev. **D53**, 2442 (1996) [hep-ph/9501407];  
J. Hisano and D. Nomura, Phys. Rev. **D59**, 116005 (1999) [hep-ph/9810479].
  122. T. Moroi, JHEP **0003**, 019 (2000) [hep-ph/0002208];  
D. Chang, A. Masiero and H. Murayama, Phys. Rev. **D67**, 075013 (2003) [hep-ph/0205111].
  123. S. Dimopoulos and L.J. Hall, Phys. Lett. **B344**, 185 (1995) [hep-ph/9411273].
  124. J. Hisano *et al.*, Phys. Lett. **B604**, 216 (2004) [hep-ph/0407169].
  125. G.L. Fogli *et al.*, Phys. Rev. **D84**, 053007 (2011) [arXiv:1106.6028].
  126. T. Schwetz, M. Tortola, and J.W.F. Valle, New J. Phys. **13**, 109401 (2011) [arXiv:1108.1376].
  127. C.H. Albright and M.-C. Chen, Phys. Rev. **D74**, 113006 (2006) [hep-ph/0608137].
  128. M. Magg and C. Wetterich, Phys. Lett. **B94**, 61 (1980);  
J. Schechter and J.W.F. Valle, Phys. Rev. **D22**, 2227 (1980);  
G. Lazarides, Q. Shafi and C. Wetterich, Nucl. Phys. **B181**, 287 (1981);  
R.N. Mohapatra and G. Senjanovic, Phys. Rev. **D23**, 165 (1981);  
G.B. Gelmini and M. Roncadelli, Phys. Lett. **B99**, 411 (1981).
  129. R. Foot, *et al.*, Z. Phys. **C44**, 441 (1989).
  130. G. 't Hooft, Nucl. Phys. **B79**, 276 (1974) A.M. Polyakov, Pis'ma Zh. Eksp. Teor. Fiz. **20**, 430 (1974) [Sov. Phys. JETP Lett. **20**, 194 (1974)];  
For a pedagogical introduction, see S. Coleman, in *Aspects of Symmetry*, Selected Erice Lectures, Cambridge University Press, Cambridge, (1985), and P. Goddard and D. Olive, Rept. on Prog. in Phys. **41**, 1357 (1978).
  131. I. De Mitri, (MACRO Collab.), Nucl. Phys. (Proc. Supp.) **B95**, 82 (2001) [arXiv:1105.5587v1].
  132. For a review, see A.D. Linde, *Particle Physics and Inflationary Cosmology*, Harwood Academic, Switzerland (1990).
  133. G.R. Dvali, H. Liu and T. Vachaspati, Phys. Rev. Lett. **80**, 2281 (1998) [hep-ph/9710301].
  134. P. Langacker and S.Y. Pi, Phys. Rev. Lett. **45**, 1 (1980).
  135. G.R. Dvali, A. Melfo and G. Senjanovic, Phys. Rev. Lett. **75**, 4559 (1995) [hep-ph/9507230].
  136. V. Rubakov, Nucl. Phys. **B203**, 311 (1982) and Institute of Nuclear Research Report No. P-0211, Moscow (1981), unpublished;  
C. Callan, Phys. Rev. **D26**, 2058 (1982);  
F. Wilczek, Phys. Rev. Lett. **48**, 1146 (1982);  
See also, S. Dawson and A.N. Schellekens, Phys. Rev. **D27**, 2119 (1983).
  137. K. Freese, M. S. Turner and D. N. Schramm, Phys. Rev. Lett. **51**, 1625 (1983).
  138. P.A.R. Ade *et al.* [BICEP2 and Planck Collabs.], Phys. Rev. Lett. **114**, 101301 (2015) [arXiv:1502.00612 [astro-ph.CO]];  
P.A.R. Ade *et al.* [Planck Collab.], arXiv:1502.02114 [astro-ph.CO].
  139. N.G. Deshpande, OITS-107;  
C.Q. Geng and R.E. Marshak, Phys. Rev. **D39**, 693 (1989);  
A. Font, L.E. Ibanez and F. Quevedo, Phys. Lett. **B228**, 79 (1989);  
K.S. Babu and R.N. Mohapatra, Phys. Rev. Lett. **63**, 938 (1989).
  140. R. Foot *et al.*, Phys. Rev. **D39**, 3411 (1989).
  141. M. Nowakowski and A. Pilaftsis, Phys. Rev. **D48**, 259 (1993) [hep-ph/9304312].
  142. T.P. T. Dijkstra, L.R. Huiszoon and A.N. Schellekens, Phys. Lett. **B609**, 408 (2005) [hep-th/0403196];  
F. Gmeiner *et al.*, JHEP **0601**, 004 (2006) [hep-th/0510170];  
B. Gato-Rivera and A.N. Schellekens, Nucl. Phys. **B883**, 529 (2014) [arXiv:1401.1782].
  143. A.Y. Ignatiev *et al.*, Phys. Lett. **B76**, 436 (1978);  
M. Yoshimura, Phys. Rev. Lett. **41**, 281 (1978) [Phys. Rev. Lett. **42**, 746 (1979)].
  144. D. Toussaint *et al.*, Phys. Rev. **D19**, 1036 (1979);  
S. Weinberg, Phys. Rev. Lett. **42**, 850 (1979);  
M. Yoshimura, Phys. Lett. **B88**, 294 (1979);  
S.M. Barr, G. Segre and H.A. Weldon, Phys. Rev. **D20**, 2494 (1979);  
D.V. Nanopoulos and S. Weinberg, Phys. Rev. **D20**, 2484 (1979);  
A. Yildiz and P.H. Cox, Phys. Rev. **D21**, 906 (1980).
  145. M. Fukugita and T. Yanagida, Phys. Rev. Lett. **89**, 131602 (2002) [hep-ph/0203194].
  146. M. Fukugita and T. Yanagida, Phys. Lett. **B174**, 45 (1986).
  147. S. Davidson and A. Ibarra, Phys. Lett. **B535**, 25 (2002) [hep-ph/0202239];  
K. Hamaguchi, H. Murayama and T. Yanagida, Phys. Rev. **D65**, 043512 (2002) [hep-ph/0109030].
  148. M. Kawasaki *et al.*, Phys. Rev. **D78**, 065011 (2008) [arXiv:0804.3745].
  149. W. Buchmuller, R.D. Peccei and T. Yanagida, Ann. Rev. Nucl. and Part. Sci. **55**, 311 (2005) [hep-ph/0502169];  
C.S. Fong, E. Nardi and A. Riotto, Adv. High Energy Phys. **2012**, 158303 (2012) [arXiv:1301.3062].
  150. G.G. Ross, "Grand Unified Theories", Benjamin/Cummings, 1984.;  
K.R. Dienes, Phys. Reports **287**, 447 (1997) [hep-th/9602045];  
P. Nath and P. Fileviez Perez, Phys. Reports **441**, 191 (2007) [hep-ph/0601023];  
S. Raby, Rept. on Prog. in Phys. **74**, 036901 (2011) [arXiv:1101.2457].

## 17. HEAVY-QUARK AND SOFT-COLLINEAR EFFECTIVE THEORY

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### 17.1. Effective Field Theories

Quantum field theories provide the most precise computational tools for describing physics at the highest energies. One of their characteristic features is that they almost inevitably involve multiple length scales. When trying to determine the value of an observable, quantum field theory demands that all possible virtual states and hence all particles be included in the calculation. Since these particles have widely different masses, the final prediction is sensitive to many scales. This fact represents a formidable challenge from a practical point of view. No realistic quantum field theories can be solved exactly, so that one needs to resort to approximation schemes; these, however, are typically most straightforward when only a single scale is involved at a time.

Effective field theories (EFTs) provide a general theoretical framework to deal with the multi-scale problems of realistic quantum field theories. This framework aims at reducing such problems to a combination of separate and simpler single-scale problems; simultaneously, however, it provides an organization scheme whereby the other scales are not omitted but allowed to play their role in a separate step of the computation. The philosophy and basic principles of this approach are very generic, and correspondingly EFTs represent a widely used method in many different areas of high-energy physics, from the low-energy scales of atomic and nuclear physics to the high-energy scales of (partly yet unknown) elementary-particle physics, see [1–3] for some early references. EFTs can play a role both within analytic perturbative computations and in the context of non-perturbative numerical simulations; One of the simplest applications of EFTs to particle physics concerns the description of an underlying theory that is only probed at energy scales  $E < \Lambda$ . Any particle with mass  $m > \Lambda$  cannot be produced as a real state and therefore only leads to short-distance virtual effects. Thus, one can construct an effective theory in which the quantum fluctuations of such heavy particles are “integrated out” from the generating functional for Green functions. This results in a simpler theory containing only those degrees of freedom that are relevant to the energy scales under consideration. In fact, the standard model of particle physics itself is widely viewed as an EFT of some yet unknown, more fundamental theory.

The development of any effective theory starts by identifying the degrees of freedom that are relevant to describe the physics at a given energy (or length) scale and constructing the Lagrangian describing the interactions among these fields. Short-distance quantum fluctuations associated with much smaller length scales are absorbed into the coefficients of the various operators in the effective theory. These coefficients are determined in a matching procedure, by requiring that the EFT reproduces the matrix elements of the full theory up to power corrections. In many cases the effective Lagrangian exhibits enhanced symmetries compared with the fundamental theory, allowing for simple and sometimes striking predictions relating different observables.

### 17.2. Heavy-Quark Effective Theory

Heavy-quark systems provide prime examples for applications of the EFT technology, because the hierarchy  $m_Q \gg \Lambda_{\text{QCD}}$  (with  $Q = b, c$ ) provides a natural separation of scales. Physics at the scale  $m_Q$  is of a short-distance nature and can be treated perturbatively, while for heavy-quark systems there is always also some hadronic physics governed by the confinement scale  $\Lambda_{\text{QCD}}$  of the strong interaction. Being able to separate the short-distance and long-distance effects associated with these two scales is crucial for any quantitative description. For instance, if the long-distance hadronic matrix elements are obtained from lattice QCD, then it is necessary to analytically compute the effects of short-wavelength modes that do not fit on the lattice. In many other instances, the long-distance physics can be encoded in a small number of hadronic parameters.

#### 17.2.1. General idea and derivation of the effective

**Lagrangian** : The simplest effective theory for heavy-quark systems is the heavy-quark effective theory (HQET) [4–7] (see [8,9] for detailed discussions). It provides a simplified description of the soft interactions of a single heavy quark with light partons. This includes the interactions that bind the heavy quark with other light partons inside heavy mesons and baryons.

A softly interacting heavy quark is nearly on-shell. Its momentum may be decomposed as  $p_Q = m_Q v + k$ , where  $v$  is the 4-velocity of the hadron containing the heavy quark. The “residual momentum”  $k$  results from the soft interactions of the heavy quark with its environment and satisfies  $v \cdot k \sim \Lambda_{\text{QCD}}$  and  $k^2 \sim \Lambda_{\text{QCD}}^2$ , which in the rest frame of the heavy hadron reduces to  $k^\mu \sim \Lambda_{\text{QCD}}$ . In the limit  $m_Q \gg \Lambda_{\text{QCD}}$ , the soft interactions do not change the 4-velocity of the heavy quark, which is therefore a conserved quantum number that is often used as a label on the effective heavy-quark fields. A nearly on-shell Dirac spinor has two large and two small components. We define

$$Q(x) = e^{-im_Q v \cdot x} [h_v(x) + H_v(x)], \quad (17.1)$$

where

$$h_v(x) = e^{im_Q v \cdot x} \frac{1 + \not{v}}{2} Q(x), \quad H_v(x) = e^{im_Q v \cdot x} \frac{1 - \not{v}}{2} Q(x) \quad (17.2)$$

are the large (“upper”) and small (“lower”) components of the spinor field, respectively. The extraction of the phase factor in Eq. (17.1) implies that the fields  $h_v$  and  $H_v$  carry the residual momentum  $k$ . The field  $H_v$  is  $1/m_Q$  suppressed relative to  $h_v$  and describes quantum fluctuations far off the mass shell. Integrating it out using its equations of motion yields the HQET Lagrangian

$$\mathcal{L}_{\text{HQET}} = \bar{h}_v i v \cdot D_s h_v + \frac{1}{2m_Q} \left[ \bar{h}_v (iD_s)^2 h_v + C_{\text{mag}}(\mu) \frac{g}{2} \bar{h}_v \sigma_{\mu\nu} G_s^{\mu\nu} h_v \right] + \dots \quad (17.3)$$

The covariant derivative  $iD_s^\mu = i\partial^\mu + gA_s^\mu$  and the field strength  $G_s^{\mu\nu}$  contain only the soft gluon field. Hard gluons have been integrated out, and their effects are contained in the Wilson coefficients of the operators in the effective Lagrangian. From the leading operator one derives the Feynman rules of HQET. The new operators entering at subleading order are referred to as the “kinetic energy” and “chromo-magnetic interaction”. The kinetic-energy operator corresponds to the first correction term in the Taylor expansion of the relativistic energy  $E = m_Q + \vec{p}^2/2m_Q + \dots$ . Lorentz invariance, which is encoded as a reparametrization invariance of the effective Lagrangian [10], ensures that its Wilson coefficient is not renormalized ( $C_{\text{kin}} \equiv 1$ ). The coefficient  $C_{\text{mag}}$  of the chromo-magnetic operator receives corrections starting at one-loop order.

**17.2.2. Spin-flavor symmetry** : The leading term in the HQET Lagrangian exhibits a global spin-flavor symmetry. Its physical meaning is that, in the infinite mass limit, the properties of hadronic systems containing a single heavy quark are insensitive to the spin and flavor of the heavy quark [11,12]. The spin symmetry results from the fact that there are no Dirac matrices in the leading term of the effective Lagrangian in Eq. (17.3), implying that the interactions of the heavy quark with soft gluons leave its spin unchanged. The flavor symmetry arises since the mass of the heavy quark does not appear at leading order. For  $n_Q$  heavy quarks moving at the same velocity, one can simply extend Eq. (17.3) by summing over  $n_Q$  identical terms for heavy-quark fields  $h_v^i$ . The result is invariant under rotations in flavor space. When combined with the spin symmetry, the symmetry group becomes promoted to  $\text{SU}(2n_Q)$ . These symmetries are broken by the operators at subleading power in the  $1/m_Q$  expansion.

The spin-flavor symmetry leads to many interesting relations between the properties of hadrons containing a heavy quark. The most direct consequences concern the spectroscopy of such states [13]. In the heavy-quark limit, the spin of the heavy quark and the total angular momentum  $j$  of the light degrees of freedom are separately conserved by the strong interactions. Because of heavy-quark symmetry, the dynamics is independent of the spin and mass of the heavy quark. Hadronic states can thus be classified by the quantum numbers (flavor, spin, parity, etc.) of the light degrees of freedom. The spin symmetry predicts that, for fixed  $j \neq 0$ , there is a

doublet of degenerate states with total spin  $J = j \pm 1/2$ . The flavor symmetry relates the properties of states with different heavy-quark flavor.

**17.2.3. Weak decay form factors :** Of particular interest are the relations between the weak decay form factors of heavy mesons, which parametrize hadronic matrix elements of currents between two mesons containing a heavy quark. These relations have been derived by Isgur and Wise [12], generalizing ideas developed by Nussinov and Wetzel [14] and Voloshin and Shifman [15]. For the purpose of this discussion, it is convenient to work with a mass-independent normalization of meson states and use velocity rather than momentum variables.

Consider the elastic scattering of a pseudoscalar meson,  $P(v) \rightarrow P(v')$ , induced by an external vector current coupled to the heavy quark contained in  $P$ , which acts as a color source moving with the meson's velocity  $v$ . The action of the current is to replace instantaneously the color source by one moving at velocity  $v'$ . Soft gluons need to be exchanged in order to rearrange the light degrees of freedom and build up the final state meson moving at velocity  $v'$ . This rearrangement leads to a form-factor suppression. The important observation is that, in the  $m_Q \rightarrow \infty$  limit, the form factor can only depend on the Lorentz boost  $\gamma = v \cdot v'$  connecting the rest frames of the initial and final-state mesons (as long as  $\gamma = \mathcal{O}(1)$ ). In the effective theory the hadronic matrix element describing the scattering process can therefore be written as

$$\langle P(v') | \bar{h}_{v'} \gamma^\mu h_v | P(v) \rangle = \xi(v \cdot v') (v + v')^\mu, \quad (17.4)$$

with a form factor  $\xi(v \cdot v')$  that is real and independent of  $m_Q$ . By flavor symmetry, the form factor remains identical when one replaces the heavy quark  $Q$  in one of the meson states by a heavy quark  $Q'$  of a different flavor, thereby turning  $P$  into another pseudoscalar meson  $P'$ . At the same time, the current becomes a flavor-changing vector current. This universal form factor is called the Isgur-Wise function [12]. For equal velocities the vector current  $J^\mu = \bar{h}_v \gamma^\mu h_v$  is conserved in the effective theory, irrespective of the flavor of the heavy quarks. The corresponding conserved charges are the generators of the flavor symmetry. It follows that the Isgur-Wise function is normalized at the point of equal velocities:  $\xi(1) = 1$ . Since the recoil energy of the daughter meson  $P'$  in the rest frame of the parent meson  $P$  is  $E_{\text{recoil}} = m_{P'}(v \cdot v' - 1)$ , the point  $v \cdot v' = 1$  is referred to as the zero-recoil limit. The heavy-quark spin symmetry leads to additional relations among weak decay form factors. It can be used to relate matrix elements involving vector mesons to those involving pseudoscalar mesons, which once again can be described completely in terms of the universal Isgur-Wise function.

The form factor relations imposed by heavy-quark symmetry describe the semileptonic decay processes  $\bar{B} \rightarrow D \ell \bar{\nu}$  and  $\bar{B} \rightarrow D^* \ell \bar{\nu}$  in the limit of infinite heavy-quark masses. They are model-independent consequences of QCD. The known normalization of the Isgur-Wise function at zero recoil can be used to obtain a model-independent measurement of the element  $|V_{cb}|$  of the Cabibbo-Kobayashi-Maskawa (CKM) matrix. The semileptonic decay  $\bar{B} \rightarrow D^* \ell \bar{\nu}$  is particularly well suited for this purpose [16]. Experimentally this is a very clean mode, since the reconstruction of the  $D^*$  meson mass provides a powerful rejection against background. From the theoretical point of view, it is ideal since the decay rate at zero recoil is protected by Luke's theorem against first-order power corrections in  $1/m_Q$  [17]. This is described in more detail in Section 12. Corrections to the heavy-quark symmetry relations for the  $\bar{B} \rightarrow D^{(*)}$  form factors near zero recoil can also be constrained using sum rules derived in the small-velocity limit [18,19].

**17.2.4. Decoupling transformation :** At leading order in  $1/m_Q$ , the couplings of soft gluons to heavy quarks in the effective Lagrangian Eq. (17.3) can be removed by the field redefinition  $h_v(x) = Y_v(x) \bar{h}_v^{(0)}(x)$ , where  $Y_v(x)$  is a soft Wilson line along the direction of  $v$ , extending from minus infinity to the point  $x$ . In terms of the new fields the leading-order HQET Lagrangian becomes  $\mathcal{L}_{\text{HQET}} = \bar{h}_v^{(0)} i v \cdot \partial h_v^{(0)}$ . It describes a free theory as far as the strong interactions of heavy quarks are concerned. However, the theory is

nevertheless non-trivial in the presence of external sources. Consider, e.g., the case of a weak-interaction heavy-quark current

$$\bar{h}_{v'} \gamma^\mu (1 - \gamma_5) h_v = \bar{h}_{v'}^{(0)} \gamma^\mu (1 - \gamma_5) Y_{v'}^\dagger Y_v h_v^{(0)}, \quad (17.5)$$

where  $v$  and  $v'$  are the velocities of the heavy mesons containing the heavy quarks. Unless the two velocities are equal, corresponding to the zero-recoil limit discussed above, the object  $Y_{v'}^\dagger Y_v$  is non-trivial, and hence the soft gluons do not decouple from the heavy quarks inside the current operator. One may interpret  $Y_{v'}^\dagger Y_v$  as a Wilson loop with a cusp at the point  $x$ , where the two paths parallel to the different velocity vectors intersect. The presence of the cusp leads to non-trivial ultra-violet behavior (for  $v \neq v'$ ), which is described by a cusp anomalous dimension  $\Gamma_c(v \cdot v')$  that was calculated at two-loop order in [20]. It coincides with the velocity-dependent anomalous dimension of heavy-quark currents, which was introduced in the context of HQET in [21]. The interpretation of heavy quarks as Wilson lines is a useful tool, which was put forward in one of the very first papers on the subject [4]. This technology will be useful in the study of the interactions of heavy quarks with collinear degrees of freedom discussed later in this review.

**17.2.5. Heavy-quark expansion for inclusive decays :** The theoretical description of inclusive decays of hadrons containing a heavy quark exploits two observations [22–26]: bound-state effects related to the initial state can be calculated using the heavy-quark expansion, and the fact that the final state consists of a sum over many hadronic channels eliminates the sensitivity to the properties of individual final-state hadrons. The second feature rests on the hypothesis of quark-hadron duality, i.e. the assumption that decay rates are calculable in QCD after a smearing procedure has been applied [27]. In semileptonic decays, the integration over the lepton spectrum provides a smearing over the invariant hadronic mass of the final state (global duality). For nonleptonic decays, where the total hadronic mass is fixed, the summation over many hadronic final states provides an averaging (local duality). Since global duality is a much weaker assumption, the theoretical control of inclusive semileptonic decays is on firmer footing.

Using the optical theorem, the inclusive decay width of a hadron  $H_b$  containing a  $b$  quark can be written in the form

$$\Gamma(H_b) = \frac{1}{M_{H_b}} \text{Im} \langle H_b | i \int d^4x T \{ \mathcal{H}_{\text{eff}}(x), \mathcal{H}_{\text{eff}}(0) \} | H_b \rangle. \quad (17.6)$$

The effective weak Hamiltonian for  $b$ -quark decays consists of dimension-6 four-fermion operators and dipole operators [28]. Because of the large mass of the  $b$  quark, it follows that the separation of fields in the time-ordered product in Eq. (17.6) is small, of order  $x \sim 1/m_b$ . It is thus possible to construct an operator-product expansion (OPE) for the time-ordered product, in which it is represented as a series of local operators in HQET. The leading operator  $\bar{h}_v h_v$  has a trivial matrix element. The next contributions arise at  $\mathcal{O}(1/m_b^2)$  and give rise to two parameters  $\mu_\pi^2(H_b)$  and  $\mu_G^2(H_b)$ , which are defined as the matrix elements of the heavy-quark kinetic energy and chromo-magnetic interaction inside the hadron  $H_b$ , respectively [29]. For the ground-state heavy mesons and baryons, one has  $\mu_G^2(B) = 3(m_{B^*}^2 - m_B^2)/4 \simeq 0.36 \text{ GeV}^2$  and  $\mu_G^2(\Lambda_b) = 0$ . Thus, the total inclusive decay rate of a hadron  $H_b$  can be written as [23,24]

$$\Gamma(H_b) = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} [c_1 + c_2 \frac{\mu_\pi^2(H_b)}{2m_b^2} + c_3 \frac{\mu_G^2(H_b)}{2m_b^2} + \mathcal{O}(\frac{1}{m_b^3}) + \dots], \quad (17.7)$$

where the prefactor arises from the loop integrations and is proportional to the fifth power of the  $b$ -quark mass. The coefficient functions  $c_i$  are calculable order by order in perturbation theory.

From the fully inclusive width in Eq. (17.7) one can obtain the lifetime of a heavy hadron via  $\tau(H_b) = 1/\Gamma(H_b)$ . Due to the universality of the leading term in the heavy-quark expansion, lifetime ratios such as  $\tau(B^-)/\tau(\bar{B}^0)$ ,  $\tau(\bar{B}_s^0)/\tau(\bar{B}^0)$  and  $\tau(\Lambda_b)/\tau(\bar{B}^0)$  are particularly sensitive to the hadronic parameters determining the



power corrections in the expansion. In order to understand these ratios theoretically, it is necessary to include phase-space enhanced power corrections of order  $(\Lambda_{\text{QCD}}/m_b)^3$  [30,31] as well as short-distance perturbative effects [32] in the calculation.

A formula analogous to Eq. (17.7) can be derived for differential distributions in specific inclusive decay processes, assuming that these distributions are integrated over a sufficiently large region of phase space to ensure quark-hadron duality. Important examples are the distributions in the lepton energy and the lepton invariant mass, as well as moments of the invariant hadronic mass distribution in the semileptonic processes  $\bar{B} \rightarrow X_u \ell \bar{\nu}$  and  $\bar{B} \rightarrow X_c \ell \bar{\nu}$ . A global fit of semileptonic decay distributions can be used to determine the CKM matrix elements  $|V_{ub}|$  and  $|V_{cb}|$  along with heavy-quark parameters such as the masses  $m_b$ ,  $m_c$  and the hadronic parameters  $\mu_\pi^2(B)$ ,  $\mu_G^2(B)$ . These determinations provide some of the most accurate values for these parameters [33].

**17.2.6. Shape functions and non-local power corrections :** In certain regions of phase space, in which the hadronic final state in an inclusive heavy-hadron decay is made up of light energetic partons, the local OPE for inclusive decays must be replaced by a more complicated expansion involving hadronic matrix elements of non-local light-ray operators [34,35]. Prominent examples are the radiative decay  $\bar{B} \rightarrow X_s \gamma$  for large photon energy  $E_\gamma$  near  $m_B/2$ , and the semileptonic decay  $\bar{B} \rightarrow X_u \ell \bar{\nu}$  at large lepton energy or small hadronic invariant mass. In these cases, the differential decay rates at leading order in the heavy-quark expansion can be written in the factorized form  $d\Gamma = H J \otimes S$  [36], where the hard function  $H$  and the jet function  $J$  are calculable in perturbation theory. The characteristic scales for these functions are set by  $m_b$  and  $(m_b \Lambda_{\text{QCD}})^{1/2}$ , respectively. The soft function

$$S(\omega) = \int \frac{dt}{4\pi} e^{-i\omega t} \langle \bar{B}(v) | \bar{h}_v(tn) Y_n(tn) Y_n^\dagger(0) h_v(0) | \bar{B}(v) \rangle \quad (17.8)$$

is a genuinely non-perturbative object called the shape function [34,35]. Here  $Y_n$  are soft Wilson lines along a light-like direction  $n$  aligned with the momentum of the hadronic final-state jet. The jet function and the shape function share a common variable  $\omega \sim \Lambda_{\text{QCD}}$ , and the symbol  $\otimes$  denotes a convolution in this variable.

While the hard functions are different for the decays  $\bar{B} \rightarrow X_s \gamma$  and  $\bar{B} \rightarrow X_u \ell \bar{\nu}$ , the jet and soft functions are identical at leading order in  $\Lambda_{\text{QCD}}/m_Q$ . This is particularly important for the shape function, which introduces non-perturbative physics into the theoretical predictions for the decay rates in the regions of experimental interest. The fact that both processes depend on the same non-perturbative function makes it possible to use the measured shape of the  $\bar{B} \rightarrow X_s \gamma$  photon spectrum to reduce the theoretical uncertainties in the determination of the CKM element  $|V_{ub}|$  from semileptonic decays. In higher orders of the heavy-quark expansion, an increasing number of subleading jet and soft functions are required to describe the decay distributions [37]. These have been analyzed in detail at order  $1/m_b$  [38–40]. In the case of  $\bar{B} \rightarrow X_s \gamma$ , some of these non-local effects survive in the total decay rate and give rise to irreducible hadronic uncertainties [41]. The technology for deriving the corresponding factorization theorems relies on the soft-collinear effective theory, to which we now turn.

### 17.3. Soft-Collinear Effective Theory

As discussed in the previous section, soft gluons that bind a heavy quark inside a heavy meson cannot change the virtuality of that heavy quark by a significant amount. The ratio  $\Lambda_{\text{QCD}}/m_Q$  provides the expansion parameter in HQET, which is a small parameter since  $m_Q \gg \Lambda_{\text{QCD}}$ . This obviously does not work when considering light quarks. However, if the energy  $Q$  of the quarks is large, the ratio  $\Lambda_{\text{QCD}}/Q$  provides a small parameter, which can be used to construct an effective theory. One major difference to HQET is that light energetic quarks cannot only emit soft gluons, but they can also emit collinear gluons (an energetic gluon in the same direction as the original quark), without parametrically changing their virtuality. Thus, to fully reproduce the long-distance physics of energetic quarks

requires that one includes their interactions with both soft and collinear particles. The resulting effective theory is therefore called soft-collinear effective theory (SCET) [42–44].

A single energetic particle can always be boosted to a frame where all momentum components have similar size, in which case there is no small expansion parameter. Thus the presence of energetic particles must refer to a reference frame defined by external kinematics. SCET has a wide range of applications; some examples are the production of energetic, light states in the decay of a heavy particle in its rest frame, the production of energetic jets in collider environments, and the scattering of energetic particles off a target at rest. In this brief review we will outline the main features of this effective theory and mention a few selected applications.

**17.3.1. General idea of the expansion :** Consider a quark with virtuality much less than its energy  $Q$ , moving along the direction  $\vec{n}$ . It is convenient to parameterize the momentum  $p_n$  of this particle in terms of its light-cone components, defined by  $(p_n^-, p_n^+, p_n^\perp) = (\bar{n} \cdot p_n, n \cdot p_n, p_n^\perp)$ , where  $n^\mu = (1, \vec{n})$  and  $\bar{n}^\mu = (1, -\vec{n})$  are light-like vectors, and  $n \cdot p_n^\perp = \bar{n} \cdot p_n^\perp = 0$ . The subscript  $n$  on the momentum indicates the direction of the collinear particle. In terms of these light-cone components, the virtuality satisfies  $p_n^2 = p_n^+ p_n^- + p_n^{\perp 2}$ . The individual components of the momentum obey

$$(p_n^-, p_n^+, p_n^\perp) \sim Q(1, \lambda^2, \lambda), \quad (17.9)$$

where  $\lambda^2 = p^2/Q^2$  is the expansion parameter of SCET. The virtuality of such an energetic particle remains parametrically unchanged if it interacts with energetic particles in the same direction  $n$ , or with soft particles with momentum scaling as

$$(p_s^-, p_s^+, p_s^\perp) \sim Q(\lambda^2, \lambda^2, \lambda^2). \quad (17.10)$$

SCET is constructed in such a way as to reproduce the long-distance dynamics arising from the interactions of collinear and soft degrees of freedom.

In the above power counting the transverse momenta of soft degrees of freedom scale as  $p_s^\perp \sim Q\lambda^2$ , which is much smaller than the transverse momenta  $p_c^\perp \sim Q\lambda$  of collinear fields. This theory is usually called SCET<sub>I</sub>. If the external kinematics require that the transverse momenta of both soft and collinear fields are of the same size,  $p_c^\perp \sim p_s^\perp$ , then the appropriate degrees of freedom have the scaling  $p_c \sim Q(1, \lambda^2, \lambda)$  and  $p_s \sim Q(\lambda, \lambda, \lambda)$ . This theory is usually called SCET<sub>II</sub> and is required, e.g., for exclusive hadronic decays such as  $\bar{B} \rightarrow D\pi$ , where the virtuality of both collinear and soft degrees of freedom are set by  $\Lambda_{\text{QCD}}$ , or for the description of transverse-momentum distributions at colliders.

**17.3.2. Leading-order Lagrangian :** The derivation of the SCET Lagrangian follows similar steps as described for HQET in Section 17.2.1. One begins by deriving the Lagrangian for a theory containing only a single collinear sector. Similar to HQET, one separates the full QCD field into two components,  $q_n(x) = \psi_n(x) + \Xi_n(x)$ , where (with  $n \cdot \bar{n} = 2$ )

$$\psi_n(x) = \frac{\not{n}\not{\bar{n}}}{4} q_n(x), \quad \Xi_n(x) = \frac{\not{n}\not{\bar{n}}}{4} q_n(x). \quad (17.11)$$

The degrees of freedom described by the field  $\Xi_n$  are far off shell and can therefore be eliminated using its equation of motion. This gives

$$\mathcal{L}_n = \bar{\psi}_n(x) \left[ i n \cdot D + i \not{D}^\perp \frac{1}{i \bar{n} \cdot D} i \not{D}^\perp \right] \frac{\not{n}\not{\bar{n}}}{2} \psi_n(x). \quad (17.12)$$

As a next step, one separates the large and residual momentum components by decomposing the collinear momentum into a “label” and a residual momentum,  $p^\mu = P^\mu + k^\mu$  with  $n \cdot P = 0$ . One then performs a phase redefinition on the collinear fields, such that  $\psi_n(x) = e^{iP \cdot x} \xi_n(x)$ . Derivatives acting on the fields  $\xi_n(x)$  now only pick out the residual momentum. Since unlike in HQET the label momentum in SCET is not conserved, one defines a label operator  $\mathcal{P}^\mu$  acting as  $\mathcal{P}^\mu \xi_n(x) = P^\mu \xi_n(x)$  [43], as well as a corresponding covariant label operator  $iD_n^\mu = \mathcal{P}^\mu + gA_n^\mu(x)$ . Note that at leading

order in power counting  $i\mathcal{D}_n^\mu$  does not contain the soft gluon field. This leads to the final SCET Lagrangian [43–46]

$$\mathcal{L}_n = \bar{\xi}_n(x) \left[ in \cdot D_n + gn \cdot A_s + i\mathcal{P}_n^\perp \frac{1}{i\bar{n} \cdot D_n} i\mathcal{P}_n^\perp \right] \frac{\not{n}}{2} \xi_n(x) + \dots, \quad (17.13)$$

where we have split  $in \cdot D$  into a collinear piece  $in \cdot D_n = in \cdot \partial + gn \cdot A_n$  and a soft piece  $gn \cdot A_s$ . This latter term gives rise to the only interaction between a collinear quark and soft gluons at leading power in  $\lambda$ . The ellipses represent higher-order interactions between soft and collinear particles.

The Lagrangian describing collinear fields in different light-like directions is simply given by the sum of the Lagrangians for each direction  $n$ , i.e.  $\mathcal{L} = \sum_n \mathcal{L}_n$ . The soft gluons are the same in each individual Lagrangian. An alternative way to understand the separation between large and small momentum components is to derive the Lagrangian of SCET in position space [46]. In this case no label operators are required, and the dependence on short-distance effects is contained in non-localities at short distances. An important difference between SCET and HQET is that the SCET Lagrangian is not corrected by short distance fluctuations. The physical reason is that in the construction described above no high-momentum modes have been integrated out [46]. Such hard modes arise when different collinear sectors are coupled via some external current (e.g. in jet production at  $e^+e^-$  or hadron colliders), or when collinear particles are produced in the rest frame of a decaying heavy object (such as in  $B$  decays). Short-distance effects are then incorporated in the Wilson coefficients of the external source operators.

**17.3.3. Collinear gauge invariance and Wilson lines :** An important aspect of SCET is the implementation of local gauge invariance. Because the effective field operators describe modes with certain momentum scalings, the effective Lagrangian respects only residual gauge symmetries. One of them satisfies the collinear scaling

$$(\bar{n} \cdot \partial, n \cdot \partial, \partial^\perp) U_n(x) \sim Q(1, \lambda^2, \lambda) U_n(x), \quad (17.14)$$

and one the soft scaling

$$(\bar{n} \cdot \partial, n \cdot \partial, \partial^\perp) U_s(x) \sim Q(\lambda^2, \lambda^2, \lambda^2) U_s(x). \quad (17.15)$$

The fact that collinear fields in different directions do not transform under the same gauge transformations implies that each collinear sector, containing particles with large momenta along a certain direction, has to be separately gauge invariant. This requires the introduction of collinear Wilson lines [43]

$$W_n(x) = P \exp \left[ -ig \int_{-\infty}^0 ds \bar{n} \cdot A_n(s\bar{n} + x) \right], \quad (17.16)$$

which transform under collinear gauge transformations according to  $W_n \rightarrow U_n W_n$ . Thus, the combination  $\chi_n \equiv W_n^\dagger \psi_n$  is gauge invariant. In a similar manner, one can define the gauge-invariant gluon field  $B_n^\mu = g^{-1} W_n^\dagger i D_n^\mu W_n$  [47,48]. Collinear operators in SCET are typically constructed from such gauge-invariant building blocks.

**17.3.4. Derivation of factorization theorems :** One of the important applications of SCET is to understand how to factorize cross sections involving energetic particles moving in different directions into simpler pieces that can either be calculated perturbatively or determined from data. Factorization theorems have been around for much longer than SCET; see [49] for a review. However, the effective theory allows for a conceptually simpler understanding of certain classes of factorization theorems [47], since most simplifications happen already at the level of the Lagrangian. The discussion in this section is valid to leading order in the power counting of the effective theory.

As discussed in the previous section, the Lagrangian of SCET does not involve any couplings between collinear particles moving in different directions. Soft gluons couple to collinear quarks only through the term  $\bar{\xi}_n gn \cdot A_s (\not{n}/2) \xi_n$  in the effective Lagrangian in Eq. (17.13). This coupling is similar to the coupling of soft gluons to heavy quarks in HQET, see Section 17.2.4. It can be removed by means of the field redefinition [44]

$$\psi_n(x) = Y_n(x) \psi_n^{(0)}(x), \quad A_n^a(x) = Y_n^{ab}(x) A_n^{b(0)}(x), \quad (17.17)$$

where  $Y_n$  and  $Y_n^{ab}$  live in the fundamental and adjoint representations of SU(3), respectively. This fact greatly facilitates proofs of factorization theorems in SCET. A QCD operator  $O(x)$  describing the interactions of collinear partons moving in different directions can thus be written as (omitting color indices for simplicity)

$$\langle O(x) \rangle = C_O(\mu) \langle \mathcal{C}_{n_a}^{(0)}(x) \mathcal{C}_{n_b}^{(0)}(x) \mathcal{C}_{n_1}^{(0)}(x) \dots \mathcal{C}_{n_N}^{(0)}(x) [\mathcal{Y}_{n_a} \mathcal{Y}_{n_b} \mathcal{Y}_{n_1} \dots \mathcal{Y}_{n_N}](x) \rangle_\mu. \quad (17.18)$$

Here  $\mathcal{C}_{n_i}^{(0)}(x)$  denotes a gauge-invariant combination of collinear fields (either quark or gluon fields) in the direction  $n_i$ . The hard matching coefficient  $C_O$  accounts for short-distance effects at the scale  $Q$ . The soft Wilson lines can either be in a color triplet or color octet representation, and are collectively denoted by  $\mathcal{Y}_{n_i}$ . Both the matrix elements and the coefficient  $C_O$  depend on the renormalization scale  $\mu$ .

Having defined the operator mediating a given process, one can calculate the cross section by squaring the operator, taking the forward matrix element and integrating over the phase space of all final-state particles. The absence of interactions between collinear degrees of freedom moving along different directions or soft degrees of freedom implies that the forward matrix element can be factorized as

$$\begin{aligned} \langle \text{in} | O(x) O^\dagger(0) | \text{in} \rangle &= |C_0(\mu)|^2 \\ &\times \langle \text{in}_a | \mathcal{C}_{n_a}(x) \mathcal{C}_{n_a}^\dagger(0) | \text{in}_a \rangle_\mu \langle \text{in}_b | \mathcal{C}_{n_b}(x) \mathcal{C}_{n_b}^\dagger(0) | \text{in}_b \rangle_\mu \\ &\times \langle 0 | \mathcal{C}_{n_1}(x) \mathcal{C}_{n_1}^\dagger(0) | 0 \rangle_\mu \dots \langle 0 | \mathcal{C}_{n_N}(x) \mathcal{C}_{n_N}^\dagger(0) | 0 \rangle_\mu \\ &\times \langle 0 | [\mathcal{Y}_{n_a} \dots \mathcal{Y}_{n_N}](x) [\mathcal{Y}_{n_a} \dots \mathcal{Y}_{n_N}]^\dagger(0) | 0 \rangle_\mu. \end{aligned} \quad (17.19)$$

Thus, the matrix element can be written as a product of simpler structures, each of which can be evaluated separately.

The vacuum matrix elements of the outgoing collinear fields are determined by jet functions  $J_i(\mu)$ . As long as the relevant scale (for example the jet mass) is sufficiently large, these functions can be calculated perturbatively. The matrix elements of the incoming collinear fields are non-perturbative objects  $B_{p/N}(\mu)$  called beam functions for parton  $p$  in nucleon  $N$  [50]. For many applications they can be related perturbatively to the well-known parton distribution functions. Finally, the vacuum matrix element of the soft Wilson lines defines a so-called soft function  $S_{ab\dots N}(\mu)$ . The shared dependence on  $x$  in the above equation implies that in momentum space the various components of the factorization theorem are convoluted with one another. Deriving this convolution requires a careful treatment of the phase-space integration, in particular treating the large and residual components of each momentum appropriately.

Putting all information together, the differential cross section for a proton-proton collision with  $N$  jet-like objects can schematically be written as

$$d\sigma \sim \sum_{ab} H_{ab}(\mu) [B_{a/P}(\mu) B_{b/P}(\mu)] \otimes [J_1(\mu) \dots J_N(\mu)] \otimes S_{ab\dots N}(\mu). \quad (17.20)$$

The hard function is equal to the square of the matching coefficient,  $H_{ab}(\mu) = |C_O(\mu)|^2$ . It should be mentioned that the most difficult part of traditional factorization proofs involves showing that so-called Glauber gluons do not spoil the above factorization theorem [51]. This question has not yet been fully addressed in the context of SCET.

**17.3.5. Resummation of large logarithms :** SCET can be used to sum the large logarithms arising in perturbative calculations to all orders in the strong coupling constant  $\alpha_s$ . In general, perturbation theory will generate a logarithmic dependence on any ratio of scales  $r$  in a problem. For processes that involve initial or final states with energy much in excess of their mass, there are two powers of logarithms for every power of  $\alpha_s$ . These are referred to as Sudakov logarithms. For widely separated scales these large logarithms can spoil the convergence of fixed-order perturbation theory. One thus needs to reorganize the expansion in such a way that  $\alpha_s L = \mathcal{O}(1)$  is kept fixed, with  $L = \ln r$ . More precisely, a proper resummation requires summing

logarithms of the form  $\alpha_s^n L^m$  with  $m \leq n+1$  in the logarithm of a cross section, by writing  $\ln \sigma \sim L g_0(\alpha_s L) + g_1(\alpha_s L) + \alpha_s g_2(\alpha_s L) + \dots$ , with functions  $g_n(x)$  that need to be determined.

The important ingredient in achieving this resummation is the fact that SCET factorizes a given cross section into simpler pieces, each of which depends on a single physical scale. The only dependence on that scale can arise through logarithms of its ratio with the renormalization scale  $\mu$ . Thus, for each of the components in the factorization theorem one can choose a renormalization scale  $\mu$  for which the large logarithmic terms are absent. Of course, the factorization formula requires a common renormalization scale  $\mu$  in all its components, and one therefore has to use the renormalization group (RG) to evolve the various component functions from their preferred scale to the common scale  $\mu$ . A novel feature of RG equations in SCET, as opposed to other EFTs, is that the anomalous dimensions entering the evolution equations of the hard, beam, jet and soft functions in a factorization formula such as Eq. (17.20) contain a single power of the logarithm of the relevant energy scale. For example, the anomalous dimension  $\gamma_H$  of the hard function has the form

$$\gamma_H(\mu) = c_H \Gamma_{\text{cusp}}(\alpha_s) \ln \frac{Q^2}{\mu^2} + \gamma(\alpha_s), \quad (17.21)$$

where  $c_H$  is a process-dependent coefficient and  $\Gamma_{\text{cusp}}$  denotes the so-called cusp anomalous dimension [20,52]. The non-cusp part  $\gamma$  of the anomalous dimension is process dependent. The presence of a logarithm in the anomalous dimension is characteristic of Sudakov problems and arises since the perturbative series contains double logarithms of scale ratios.

The anomalous dimension  $\gamma_H$  is known at two-loop order for arbitrary  $n$ -parton amplitudes containing massless or massive external partons [53–56]. Solving the RG equations one can systematically resum all large logarithms of scale ratios in the factorized cross section and express the functions  $g_n(\alpha_s L)$  introduced above in terms of ratios of running coupling constants. In order to compute the first two terms  $L g_0(\alpha_s L) + g_1(\alpha_s L)$  in  $\ln \sigma$ , corresponding to the next-to-leading logarithmic (NLL) approximation, one needs two-loop expressions for the cusp anomalous dimension and  $\beta$  function, one-loop expressions for the non-cusp pieces in the anomalous dimensions, and tree-level matching conditions for all component functions at their characteristic scales. To calculate the next term  $\alpha_s g_2(\alpha_s L)$  in the expansion, corresponding to NNLL order, one needs to go one order higher in the loop expansion, and so on.

**17.3.6. Factorization and resummation in SCET<sub>II</sub>:** The effective theory SCET<sub>II</sub> contains collinear and soft particles with momenta scaling as  $(p_n^-, p_n^+, p_n^\perp) \sim Q(1, \lambda^2, \lambda)$  and  $(p_s^-, p_s^+, p_s^\perp) \sim Q(\lambda, \lambda, \lambda)$ . They have the same small virtuality ( $p_n^2 \sim p_s^2 \sim Q^2 \lambda^2$ ) but differ in their rapidities. An important class of observables, for which this scaling is relevant, contains cross sections for processes in which the transverse momenta of particles are constrained by external kinematics. The prime example are the transverse-momentum distributions of electroweak gauge bosons or Higgs bosons produced at hadron colliders. The parton transverse momenta are constrained by the fact that their vector sum must be equal and opposite to the transverse momentum  $q_T$  of the boson. Standard RG evolution in the effective theory controls the logarithms arising from the fact that the virtualities of the collinear and soft modes are much smaller than the hard scale  $Q$  in the process (the boson mass). However, additional large logarithms arise since the rapidities of collinear and soft modes are parametrically different, such that  $e^{|\ln y_c - y_s|} \sim 1/\lambda$ . These logarithms need to be factorized in the cross section and resummed by other means.

Two equivalent approaches exist for how to deal with the additional rapidity logarithms. In the first approach, they are interpreted as a consequence of a “collinear anomaly” of the effective theory SCET<sub>II</sub>, resulting from the fact that a classical rescaling symmetry of the effective Lagrangian is broken by quantum effects [57]. The extra large logarithms can be resummed by means of simple differential equations, which typically state that (in an appropriate space) the logarithm of the cross section contains only a single logarithm of  $\lambda \sim q_T/Q$ , to all orders in perturbation theory. An alternative approach to resum the rapidity logarithms uses the “rapidity

renormalization group”, in which the relevant differential equations are obtained by considering a new type of scale variation in a parameter  $\nu$ , which separates the phase space for collinear and soft particles along a hyperbola in the  $(p_-, p_+)$  plane [58]. In contrast to the standard RG, there is no running coupling involved in the  $\nu$  evolution, since the different contributions live at the same virtuality.

SCET<sub>II</sub> also plays an important role in the study of factorization for a variety of exclusive  $B$  meson decays, such as  $\bar{B} \rightarrow \pi \ell \nu$ ,  $\bar{B} \rightarrow K^* \gamma$  and  $B \rightarrow \pi \pi$ , for which the virtualities of energetic (collinear) final-state particles are of order  $\Lambda_{\text{QCD}}$ , which is also the scale for the soft light degrees of freedom contained in the initial-state  $B$  meson.

**17.3.7. Applications:** Most of the applications of SCET are either in flavor physics, where the decay of a heavy  $B$  meson can give rise to energetic light partons, or in collider physics, where the presence of jets naturally leads to collimated sets of energetic particles. For many of these applications alternative approaches existed before the invention of SCET, but the effective theory has opened up alternative ways to understand the physics of these processes. For several examples, however, SCET has allowed new insights. The investigation of heavy-to-light form factors has been instrumental for understanding factorization in exclusive semileptonic  $B$  decays [59]. SCET has also provided a field-theoretic basis for the QCD factorization approach to exclusive, non-leptonic decays of  $B$  mesons [60]. Using SCET methods, proofs of factorization were derived for the color-allowed decay  $\bar{B}^0 \rightarrow D^+ \pi^-$  [61], the color-suppressed decay  $\bar{B}^0 \rightarrow D^0 \pi^0$  [62], and the radiative decay  $\bar{B} \rightarrow K^* \gamma$  [63]. Further examples are factorization theorems and the resummation of endpoint logarithms for quarkonia production [64], the resummation of large logarithmic terms for the thrust [65] and jet broadening [66] distributions in  $e^+e^-$  annihilation beyond NLL order, the development of new factorizable observables to veto extra jets [67], all-orders factorization theorems for processes containing electroweak Sudakov logarithms [68], and the resummation of threshold (soft gluon) logarithms for several important processes at hadron colliders [69–71]. Recently, there has been a lot of activity describing  $p_T$ -based resummation at hadron colliders. Examples are the transverse-momentum distributions of electroweak bosons [57] and jets [72]. We now describe three applications in more detail.

Event-shape distributions, in particular the thrust distribution, have been measured to high accuracy at LEP [73]. They can be used for a determination of the strong coupling constant  $\alpha_s$ . SCET has increased the theoretical accuracy in the calculations of the thrust and C-parameter distributions significantly. First, it has allowed to increase the perturbative accuracy of the thrust spectrum. The resummation of logarithms of  $\tau$ , which become important for  $\tau \ll 1$ , has been performed to N<sup>3</sup>LL [65], two orders beyond what was previously available. Combining this resummation with the known two-loop spectrum [74,75] gives precise perturbative predictions both at small and large values of  $\tau$ . Second, the factorization of the cross section in SCET has made it possible to include non-perturbative physics through a shape function, in analogy with the  $B$ -physics case discussed in Section 17.2.6. Comparing the theoretical predictions to the measured thrust and C-parameter distributions yields a precise value of the strong coupling constant  $\alpha_s(m_Z)$ , which however is lower than the average value cited in Chap. 9 [76] by several standard deviations [77,78].

The Higgs-boson production cross section in gluon fusion at the LHC, defined with a jet veto stating that no jet in the final state has transverse momentum above a threshold  $p_T^{\text{veto}}$ , can be factorized in the form [79,80] (see [81] for a corresponding calculation outside the SCET framework)

$$\begin{aligned} \sigma(p_T^{\text{veto}}) &= H(m_H, \mu) \left( \frac{\nu_B}{\nu_S} \right)^{-2F_{gg}(R, p_T^{\text{veto}}, \mu)} S_{gg}(R, p_T^{\text{veto}}, \mu, \frac{\nu_S}{p_T^{\text{veto}}}) \\ &\times \int_\tau^1 \frac{dz}{z} B_{g/P}(z, R, p_T^{\text{veto}}, \mu, \frac{\nu_B}{m_H}) B_{g/P}(\frac{\tau}{z}, R, p_T^{\text{veto}}, \mu, \frac{\nu_B}{m_H}), \end{aligned} \quad (17.22)$$

where  $\tau = m_H^2/s$ , and  $\mu \sim p_T^{\text{veto}}$  is a common factorization scale. The beam functions  $B_{g/P}$ , the soft function  $S_{gg}$  and the exponent  $F_{gg}$  all depend on the jet radius  $R$  as well as the jet clustering algorithm.

The scale dependence of the hard function  $H$  is controlled by standard RG evolution in SCET. The beam functions can be factorized further into calculable collinear kernels convoluted with parton distribution functions. In addition to the renormalization scale  $\mu$ , the beam and soft functions depend on two rapidity scales  $\nu_B \sim m_H$  and  $\nu_S \sim p_T^{\text{veto}}$ , respectively. In [79] the default values  $\nu_B = m_H$  and  $\nu_S = p_T^{\text{veto}}$  are used for these scales, and the soft function  $S_{gg}$  is absorbed into the beam functions. In [80] the exponent  $F_{gg}$  is called  $-\gamma_\nu^g/2$ . The second factor on the right-hand side of the factorization formula Eq. (17.22), which resums large rapidity logarithms, implies that the logarithm of the jet-veto cross section contains a single large logarithm  $\ln \sigma = -2F_{gg}(R, p_T^{\text{veto}}, \mu) \ln(m_H/p_T^{\text{veto}}) + \dots$  not contained in the hard function. Its coefficient can be calculated in fixed-order perturbation theory.

Obtaining more precise fixed-order calculations has been an important goal for many years. A major difficulty in these calculations is the proper handling of the infrared singularities that arise in both virtual and real contributions. Recently, a proposal has been made to use a so-called  $N$ -jettiness ( $\mathcal{T}_N$ ) subtraction/slicing method to obtain the NNLO result from a much easier NLO calculation, combined with information about the singular dependence of the cross section on the  $\mathcal{T}_N$  resolution variable [82,83]. While the NLO calculations can be performed using well established techniques, the singular dependence on  $\mathcal{T}_N$  can be calculated using SCET at NNLO.

#### 17.4. Open issues and perspectives

HQET has successfully passed many experimental tests, and there are not many open questions that still need to be addressed. One concept that has not been derived from first principles is the notion of quark-hadron duality, which underlies the application of HQET to the description of inclusive decays of  $B$  mesons. The validity of global duality (at energies even lower than those relevant in  $B$  decays) has been tested experimentally using high-precision data on semileptonic  $B$  decays and on hadronic  $\tau$  decays, and good agreement between theory and data was found. However, assigning a theoretical uncertainty due to possible duality violations remains a difficult task. Another known issue is that the measured values of the CKM elements  $|V_{cb}|$  and  $|V_{ub}|$  extracted from exclusive or inclusive decays of  $B$  mesons differ from each other by several standard deviations (see Ref. 84). Both measurements rely on the heavy-quark limit, and the uncertainties quoted include theoretical estimates of the effects of power corrections arising from the finite  $b$ -quark mass. It remains an open question whether the discrepancies are due to underestimated theoretical or experimental uncertainties, or whether they may hint to the existence of new physics.

SCET, on the other hand, is still an active field of research, and new results are being obtained regularly. An active area of research is the understanding of non-global logarithms arising in hadron-collider processes with jets [85,86]. SCET-based fixed-order calculations have helped to shed some light on the nature of these logarithms [87–89]. Another active field concerns the study of Glauber gluons in SCET [90] and their relation to the BFKL equation familiar from small- $x$  physics [91]. A solid understanding of these issues will be necessary to make factorization proofs in SCET more rigorous. Glauber gluons also play an important role in SCET-based analysis of jet propagation in dense QCD media [92–95], which gives rise to the jet-quenching phenomenon in heavy-ion collisions. An important open questions facing some applications of SCET concerns factorized expressions containing endpoint-divergent convolution integrals. This problem arises, for example, in the description of heavy-to-light form factors such as  $F_{B \rightarrow \pi}(q^2)$  at large recoil [96].

We close this short review by mentioning a particularly nice application combining the methods of heavy-particle EFTs such as HQET and non-relativistic QCD with SCET in the context of describing the interactions of heavy dark matter (with mass  $M \gg v$ ) with SM particles. In [97] it was realized that the interactions of heavy, weakly interacting massive particles (WIMPs) with nuclear targets can be described in a model-independent way using heavy-particle EFTs. The WIMPs are charged under  $SU(2)_L$  and can interact with electroweak gauge bosons and the Higgs boson. The WIMP EFT

was later extended by describing the produced, highly energetic electroweak gauge bosons in terms of soft or collinear fields in SCET [98–100]. This allows one to systematically separate all relevant mass scales, resum electroweak Sudakov logarithms and disentangle the so-called Sommerfeld enhancement from the short-distance hard annihilation process.

#### References:

1. E. Witten, Nucl. Phys. **B122**, 109 (1977).
2. S. Weinberg, Phys. Lett. **B91**, 51 (1980).
3. L.J. Hall, Nucl. Phys. **B178**, 75 (1981).
4. E. Eichten, B. Hill, Phys. Lett. **B234**, 511 (1990).
5. H. Georgi, Phys. Lett. **B240**, 447 (1990).
6. B. Grinstein, Nucl. Phys. **B339**, 253 (1990).
7. T. Mannel, W. Roberts, Z. Ryzak, Nucl. Phys. **B368**, 204 (1992).
8. M. Neubert, Phys. Reports **245**, 259 (1994).
9. A.V. Manohar, M.B. Wise, Camb. Monogr. Part. Phys. Nucl. Phys. Cosmol. **10**, 1 (2000).
10. M.E. Luke, A.V. Manohar, Phys. Lett. **B286**, 348 (1992).
11. E.V. Shuryak, Phys. Lett. **B93**, 134 (1980).
12. N. Isgur, M.B. Wise, Phys. Lett. **B232**, 113 (1989); Phys. Lett. **B237**, 527 (1990).
13. N. Isgur, M.B. Wise, Phys. Rev. Lett. **66**, 1130 (1991).
14. S. Nussinov, W. Wetzel, Phys. Rev. **D36**, 130 (1987).
15. M.A. Shifman, M.B. Voloshin, Sov. J. Nucl. Phys. **45**, 292 (1987); Sov. J. Nucl. Phys. **47**, 511 (1988).
16. M. Neubert, Phys. Lett. **B264**, 455 (1991).
17. M.E. Luke, Phys. Lett. **B252**, 447 (1990).
18. I.I.Y. Bigi *et al.*, Phys. Rev. **D52**, 196 (1995).
19. N. Uraltsev, Phys. Lett. **B501**, 86 (2001).
20. G.P. Korchemsky, A.V. Radyushkin, Nucl. Phys. **B283**, 342 (1987).
21. A.F. Falk *et al.*, Nucl. Phys. **B343**, 1 (1990).
22. J. Chay, H. Georgi, B. Grinstein, Nucl. Phys. **B247**, 399 (1990).
23. I.I.Y. Bigi, N.G. Uraltsev, A.I. Vainshtein, Phys. Lett. **B293**, 430 (1992); I.I.Y. Bigi *et al.*, Phys. Rev. Lett. **71**, 496 (1993).
24. A.V. Manohar, M.B. Wise, Phys. Rev. **D49**, 1310 (1994).
25. T. Mannel, Nucl. Phys. **B413**, 396 (1994).
26. A.F. Falk, M.E. Luke, M.J. Savage, Phys. Rev. **D49**, 3367 (1994).
27. E.C. Poggio, H.R. Quinn, S. Weinberg, Phys. Rev. **D13**, 1958 (1976).
28. G. Buchalla, A.J. Buras, M.E. Lautenbacher, Rev. Mod. Phys. **68**, 1125 (1996).
29. A.F. Falk, M. Neubert, Phys. Rev. **D47**, 2965 (1993).
30. M. Neubert, C.T. Sachrajda, Nucl. Phys. **B483**, 339 (1997).
31. M. Beneke, G. Buchalla, I. Dunietz, Phys. Rev. **D54**, 4419 (1996).
32. M. Beneke, G. Buchalla, C. Greub, A. Lenz and U. Nierste, Phys. Lett. **B459**, 631 (1999); Nucl. Phys. **B639**, 389 (2002).
33. C.W. Bauer *et al.*, Phys. Rev. **D70**, 094017 (2004).
34. M. Neubert, Phys. Rev. **D49**, 4623 (1994).
35. I.I.Y. Bigi *et al.*, Int. J. Mod. Phys. **A9**, 2467 (1994).
36. G.P. Korchemsky, G.F. Sterman, Phys. Lett. **B340**, 96 (1994).
37. C.W. Bauer, M.E. Luke, T. Mannel, Phys. Rev. **D68**, 094001 (2003).
38. K.S.M. Lee, I.W. Stewart, Nucl. Phys. **B721**, 325 (2005).
39. S.W. Bosch, M. Neubert, G. Paz, JHEP **0411**, 073 (2004).
40. M. Beneke *et al.*, JHEP **0506**, 071 (2005).
41. M. Benzke *et al.*, JHEP **1008**, 099 (2010).
42. C.W. Bauer, S. Fleming, M.E. Luke, Phys. Rev. **D63**, 014006 (2000); C.W. Bauer *et al.*, Phys. Rev. **D63**, 114020 (2001).
43. C.W. Bauer, I.W. Stewart, Phys. Lett. **B516**, 134 (2001).
44. C.W. Bauer, D. Pirjol, I.W. Stewart, Phys. Rev. **D65**, 054022 (2002).
45. J. Chay, C. Kim, Phys. Rev. **D65**, 114016 (2002).
46. M. Beneke *et al.*, Nucl. Phys. **B643**, 431 (2002).
47. C.W. Bauer *et al.*, Phys. Rev. **D66**, 014017 (2002).
48. R.J. Hill, M. Neubert, Nucl. Phys. **B657**, 229 (2003).

49. J.C. Collins, D.E. Soper, G.F. Sterman, *Adv. Ser. Direct. High Energy Phys.* **5**, 1 (1988).
50. I.W. Stewart, F.J. Tackmann, W.J. Waalewijn, *Phys. Rev. D* **81**, 094035 (2010).
51. J.C. Collins, D.E. Soper, G. Sterman, *Nucl. Phys. B* **261**, 104 (1985); *Nucl. Phys. B* **308**, 833 (1988).
52. I.A. Korchemskaya, G.P. Korchemsky, *Phys. Lett. B* **287**, 169 (1992).
53. T. Becher, M. Neubert, *Phys. Rev. Lett.* **102**, 162001 (2009); *JHEP* **0906**, 081 (2009).
54. E. Gardi, L. Magnea, *JHEP* **0903**, 079 (2009).
55. A. Mitov, G.F. Sterman, I. Sung, *Phys. Rev. D* **79**, 094015 (2009).
56. A. Ferroglia *et al.*, *Phys. Rev. Lett.* **103**, 201601 (2009).
57. T. Becher, M. Neubert, *Eur. Phys. J. C* **71**, 1665 (2011).
58. J.Y. Chiu *et al.*, *JHEP* **1205**, 084 (2012).
59. M. Beneke and T. Feldmann, *Nucl. Phys. B* **685**, 249 (2004).
60. M. Beneke *et al.*, *Phys. Rev. Lett.* **83**, 1914 (1999); *Nucl. Phys. B* **591**, 313 (2000).
61. C.W. Bauer, D. Pirjol, I.W. Stewart, *Phys. Rev. Lett.* **87**, 201806 (2001).
62. S. Mantry, D. Pirjol, I.W. Stewart, *Phys. Rev. D* **68**, 114009 (2003).
63. T. Becher, R.J. Hill, M. Neubert, *Phys. Rev. D* **72**, 094017 (2005).
64. S. Fleming, A.K. Leibovich, T. Mehen, *Phys. Rev. D* **68**, 094011 (2003).
65. T. Becher, M.D. Schwartz, *JHEP* **0807**, 034 (2008).
66. T. Becher, G. Bell, M. Neubert, *Phys. Lett. B* **704**, 276 (2011).
67. I.W. Stewart, F.J. Tackmann, W.J. Waalewijn, *Phys. Rev. D* **81**, 094035 (2010).
68. J.-y. Chiu, R. Kelley, A.V. Manohar, *Phys. Rev. D* **78**, 073006 (2008).
69. T. Becher, M. Neubert, G. Xu, *JHEP* **0807**, 030 (2008).
70. V. Ahrens *et al.*, *Eur. Phys. J. C* **62**, 333 (2009);  
V. Ahrens *et al.*, *JHEP* **1009**, 097 (2010).
71. X. Liu, S. Mantry, F. Petriello, *Phys. Rev. D* **86**, 074004 (2012).
72. X. Liu, F. Petriello, *Phys. Rev. D* **87**, 014018 (2013).
73. For a review, see: S. Kluth, *Rept. on Prog. in Phys.* **69**, 1771 (2006).
74. A. Gehrmann-De Ridder *et al.*, *Phys. Rev. Lett.* **99**, 132002 (2007).
75. S. Weinzierl, *Phys. Rev. Lett.* **101**, 162001 (2008).
76. S. Bethke, G. Dissertori, and G.P. Salam, “*Quantum Chromodynamics*” in this *Review*, [pdg.lbl.gov](http://pdg.lbl.gov).
77. R. Abbate *et al.*, *Phys. Rev. D* **83**, 074021 (2011).
78. A.H. Hoang *et al.*, *Phys. Rev. D* **91**, 094018 (2015).
79. T. Becher, M. Neubert, *JHEP* **1207**, 108 (2012).
80. I.W. Stewart *et al.*, *Phys. Rev. D* **89**, 054001 (2014).
81. A. Banfi *et al.*, *Phys. Rev. Lett.* **109**, 202001 (2012).
82. R. Boughezal, X. Liu, F. Petriello, *Phys. Rev. D* **91**, 094035 (2015).
83. J. Gaunt *et al.*, *JHEP* **1509**, 058 (2015).
84. R. Kowalewski, T. Mannel, “*Semileptonic bottom hadron decays and the determination of  $V_{cb}$  and  $V_{ub}$* ”, in this *Review*, [pdg.lbl.gov](http://pdg.lbl.gov).
85. M. Dasgupta, G.P. Salam, *Phys. Lett. B* **512**, 323 (2001); *JHEP* **0203**, 017 (2002).
86. R.B. Appleby, M.H. Seymour, *JHEP* **0212**, 063 (2002).
87. R. Kelley *et al.*, *Phys. Rev. D* **84**, 045022 (2011); *Phys. Rev. D* **86**, 054017 (2012).
88. A. Hornig *et al.*, *JHEP* **1108**, 054 (2011).
89. A. von Manteuffel, R.M. Schabinger, H.X. Zhu, *JHEP* **1403**, 139 (2014).
90. C.W. Bauer, B.O. Lange, G. Ovanessian, *JHEP* **1107**, 077 (2011).
91. S. Fleming, *Phys. Lett. B* **735**, 266 (2014).
92. A. Idilbi, A. Majumder, *Phys. Rev. D* **80**, 054022 (2009).
93. G. Ovanessian, I. Vitev, *JHEP* **1106**, 080 (2011).
94. M. Benzke *et al.*, *JHEP* **1302**, 129 (2013).
95. Z.B. Kang *et al.*, *Phys. Rev. Lett.* **114**, 092002 (2015).
96. M. Beneke, T. Feldmann, *Nucl. Phys. B* **592**, 3 (2001).
97. R.J. Hill, M.P. Solon, *Phys. Lett. B* **707**, 539 (2012); *Phys. Rev. Lett.* **112**, 211602 (2014).
98. M. Bauer *et al.*, *JHEP* **1501**, 099 (2015);  
R.J. Hill, M.P. Solon, *Phys. Rev. D* **91**, 043505 (2015).
99. M. Baumgart, I.Z. Rothstein, V. Vaidya, *Phys. Rev. Lett.* **114**, 211301 (2015).
100. G. Ovanessian, T.R. Slatyer, I.W. Stewart, *Phys. Rev. Lett.* **114**, 211302 (2015).

## 18. LATTICE QUANTUM CHROMODYNAMICS

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Many physical processes considered in the Review of Particle Properties (RPP) involve hadrons. The properties of hadrons—which are composed of quarks and gluons—are governed primarily by Quantum Chromodynamics (QCD) (with small corrections from Quantum Electrodynamics [QED]). Theoretical calculations of these properties require non-perturbative methods, and Lattice Quantum Chromodynamics (LQCD) is a tool to carry out such calculations. It has been successfully applied to many properties of hadrons. Most important for the RPP are the calculation of electroweak form factors, which are needed to extract Cabibbo-Kobayashi-Maskawa (CKM) matrix elements when combined with the corresponding experimental measurements. LQCD has also been used to determine other fundamental parameters of the standard model, in particular the strong coupling constant and quark masses.

This review describes the theoretical foundations of LQCD and sketches the methods used to calculate the quantities relevant for the RPP. It also describes the various sources of error that must be controlled in a LQCD calculation. Results for hadronic quantities are given in the corresponding dedicated reviews.

### 18.1. Lattice regularization of QCD

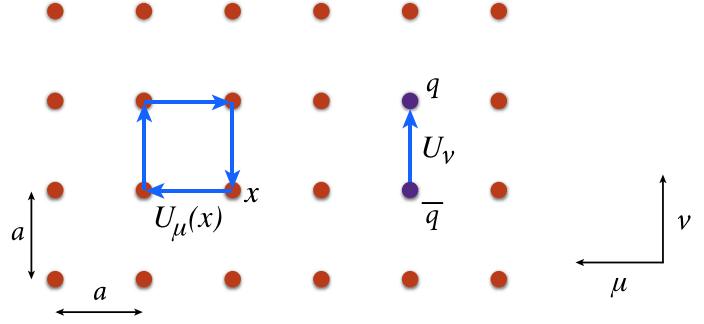
Gauge theories form the building blocks of the Standard Model. While the SU(2) and U(1) parts have weak couplings and can be studied accurately with perturbative methods, the SU(3) component—QCD—is only amenable to a perturbative treatment at high energies. The growth of the coupling constant in the infrared—the flip-side of asymptotic freedom—requires the use of non-perturbative methods to determine the low energy properties of QCD. Lattice gauge theory, proposed by K. Wilson in 1974 [1], provides such a method, for it gives a non-perturbative definition of vector-like gauge field theories like QCD. In lattice regularized QCD—commonly called lattice QCD or LQCD—Euclidean space-time is discretized, usually on a hypercubic lattice with lattice spacing  $a$ , with quark fields placed on sites and gauge fields on the links between sites. The lattice spacing plays the role of the ultraviolet regulator, rendering the quantum field theory finite. The continuum theory is recovered by taking the limit of vanishing lattice spacing, which can be reached by tuning the bare coupling constant to zero according to the renormalization group.

Unlike dimensional regularization, which is commonly used in continuum QCD calculations, the definition of LQCD does not rely on the perturbative expansion. Indeed, LQCD allows non-perturbative calculations by numerical evaluation of the path integral that defines the theory.

Practical LQCD calculations are limited by the availability of computational resources and the efficiency of algorithms. Because of this, LQCD results come with both statistical and systematic errors, the former arising from the use of Monte-Carlo integration, the latter, for example, from the use of non-zero values of  $a$ . There are also different ways in which the QCD action can be discretized, and all must give consistent results in the continuum limit,  $a \rightarrow 0$ . It is the purpose of this review to provide an outline of the methods of LQCD, with particular focus on applications to particle physics, and an overview of the various sources of error. This should allow the reader to better understand the LQCD results that are presented in other reviews, primarily those on “Quark Masses”, “Quantum Chromodynamics”, “CKM quark-mixing matrix”, “ $V_{ud}$ ,  $V_{us}$ , Cabibbo angle and CKM Unitarity” and “Semileptonic B-meson decays and the determination of  $V_{cb}$  and  $V_{ub}$ ”. For more extensive explanations the reader should consult the available textbooks or lecture notes, the most up-to-date of which are Refs. 2–4.

#### 18.1.1. Gauge invariance, gluon fields and the gluon action :

A key feature of the lattice formulation of QCD is that it preserves gauge invariance. This is in contrast to perturbative calculations, where gauge fixing is an essential step. The preservation of gauge invariance leads to considerable simplifications, e.g. restricting the form of operators that can mix under renormalization.



**Figure 18.1:** Sketch of a two-dimensional slice through the  $\mu - \nu$  plane of a lattice, showing gluon fields lying on links and forming either the plaquette product appearing in the gauge action or a component of the covariant derivative connecting quark and antiquark fields.

The gauge transformations of lattice quark fields are just as in the continuum:  $q(x) \rightarrow V(x)q(x)$  and  $\bar{q}(x) \rightarrow \bar{q}(x)V^\dagger(x)$ , with  $V(x)$  an arbitrary element of SU(3). The only difference is that the Euclidean space-time positions  $x$  are restricted to lie on the sites of the lattice, i.e.  $x = a(n_1, n_2, n_3, n_4)$  for a hypercubic lattice, with the  $n_j$  being integers. Quark bilinears involving different lattice points can be made gauge invariant by introducing the gluon field  $U_\mu(x)$ . For example, for adjacent points the bilinear is  $\bar{q}(x)U_\mu(x)q(x+a\hat{\mu})$ , with  $\hat{\mu}$  the unit vector in the  $\mu$ 'th direction. (This form is used in the construction of the lattice covariant derivative.) This is illustrated in Fig. 18.1. The gluon field (or “gauge link”) is an element of the group, SU(3), in contrast to the continuum field  $A_\mu$  which takes values in the Lie algebra. The bilinear is invariant if  $U_\mu$  transforms as  $U_\mu(x) \rightarrow V(x)U_\mu(x)V^\dagger(x+a\hat{\mu})$ . The lattice gluon field is naturally associated with the link joining  $x$  and  $x+a\hat{\mu}$ , and corresponds in the continuum to a Wilson line connecting these two points,  $P \exp(i \int_x^{x+a\hat{\mu}} dx_\mu A_\mu^{\text{cont}}(x))$  (where  $P$  indicates a path-ordered integral, and the superscript on  $A_\mu$  indicates that it is a continuum field). The trace of a product of the  $U_\mu(x)$  around any closed loop is easily seen to be gauge invariant and is the lattice version of a Wilson loop.

The simplest possible gauge action, usually called the Wilson gauge action, is given by the product of gauge links around elementary plaquettes:

$$S_g = \beta \sum_{x, \mu, \nu} \left[ 1 - \frac{1}{3} \text{ReTr}[U_\mu(x)U_\nu(x+a\hat{\mu})U_\mu^\dagger(x+a\hat{\nu})U_\nu^\dagger(x)] \right]. \quad (18.1)$$

This is illustrated in Fig. 18.1. For small  $a$ , assuming that the fields are slowly varying, one can expand the action in powers of  $a$  using  $U_\mu(x) = \exp(iaA_\mu(x))$ . Keeping only the leading non-vanishing term, and replacing the sum with an integral, one finds the continuum form,

$$S_g \longrightarrow \int d^4x \frac{1}{4g_{\text{lat}}^2} \text{Tr}[F_{\mu\nu}^2(x)], \quad (F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu + i[A_\mu, A_\nu]) \quad (18.2)$$

as long as one chooses  $\beta = 6/g_{\text{lat}}^2$  for the lattice coupling. In this expression,  $g_{\text{lat}}$  is the bare coupling constant in the lattice scheme, which can be related (by combining continuum and lattice perturbation theory) to a more conventional coupling constant such as that in the  $\overline{\text{MS}}$  scheme (see Sec. 18.3.4 below).

In practice, the lattice spacing  $a$  is non-zero, leading to discretization errors. In particular, the lattice breaks Euclidean rotational invariance (which is the Euclidean version of Lorentz invariance) down to a

discrete hypercubic subgroup. One wants to reduce discretization errors as much as possible. A very useful tool for understanding and then reducing discretization errors is the Symanzik effective action: the interactions of quarks and gluons with momenta low compared to the lattice cutoff ( $|p| \ll 1/a$ ) are described by a continuum action consisting of the standard continuum terms (e.g. the gauge action given in Eq. (18.2)) augmented by higher dimensional operators suppressed by powers of  $a$  [5]. For the Wilson lattice gauge action, the leading corrections come in at  $\mathcal{O}(a^2)$ . They take the form  $\sum_j a^2 c_j O_6^{(j)}$ , with the sum running over all dimension-six operators  $O_6^{(j)}$  allowed by the *lattice* symmetries, and  $c_j$  unknown coefficients. Some of these operators violate Euclidean rotational invariance, and all of them lead to discretization errors of the form  $a^2 \Lambda^2$ , where  $\Lambda$  is a typical momentum scale for the quantity being calculated. These errors can, however, be reduced by adding corresponding operators to the lattice action and tuning their coefficients to eliminate the dimension-six operators in the effective action to a given order in perturbation theory or even non-perturbatively. This is the idea of the Symanzik improvement program [5]. In the case of the gauge action, one adds Wilson loops involving six gauge links (as opposed to the four links needed for the original plaquette action, Eq. (18.1)) to define the  $\mathcal{O}(a^2)$  improved (or “Symanzik”) action [6]. In practical implementations, the improvement is either at tree-level (so that residual errors are proportional to  $\alpha_s a^2$ , where the coupling is evaluated at a scale  $\sim 1/a$ ), or at one loop order (errors proportional to  $\alpha_s^2 a^2$ ). Another popular choice is motivated by studies of renormalization group (RG) flow. It has the same terms as the  $\mathcal{O}(a^2)$  improved action but with different coefficients, and is called the RG-improved or “Iwasaki” action [7].

### 18.1.2. Lattice fermions :

Discretizing the fermion action turns out to involve subtle issues, and the range of actions being used is more extensive than for gauge fields. Recall that the continuum fermion action is  $S_f = \int d^4x \bar{q}[iD_\mu \gamma_\mu + m_q]q$ , where  $D_\mu = \partial_\mu + iA_\mu$  is the gauge-covariant derivative. The simplest discretization replaces the derivative with a symmetric difference:

$$D_\mu q(x) \longrightarrow \frac{1}{2a}[U_\mu(x)q(x+a\hat{\mu}) - U_\mu(x-a\hat{\mu})^\dagger q(x-a\hat{\mu})]. \quad (18.3)$$

The factors of  $U_\mu$  ensure that  $D_\mu q(x)$  transforms under gauge transformations in the same way as  $q(x)$ , so that the discretized version of  $\bar{q}(x)D_\mu \gamma_\mu q(x)$  is gauge invariant. The choice in Eq. (18.3) leads to the so-called naive fermion action. This, however, suffers from the fermion doubling problem—in  $d$  dimensions it describes  $2^d$  equivalent fermion fields in the continuum limit. The appearance of the extra “doubler” fermions is related to the deeper theoretical problem of formulating chirally symmetric fermions on the lattice. This is encapsulated by the Nielsen-Ninomiya theorem [8]: one cannot define lattice fermions having exact, continuum-like chiral symmetry without producing doublers. Naive lattice fermions do have chiral symmetry but at the cost of introducing 15 unwanted doublers (for  $d = 4$ ).

There are a number of different strategies for dealing with the doubling problem, each with their own theoretical and computational advantages and disadvantages. Wilson fermions [1] add a term proportional to  $a\bar{q}\Delta q$  to the fermion action (the “Wilson term”—in which  $\Delta$  is a covariant lattice Laplacian). This gives a mass of  $\mathcal{O}(1/a)$  to the doublers, so that they decouple in the continuum limit. The Wilson term, however, violates chiral symmetry, and also introduces discretization errors linear in  $a$ . A commonly used variant that eliminates the  $\mathcal{O}(a)$  discretization error is the  $\mathcal{O}(a)$ -improved Wilson (or “clover”) fermion [9]. In this application of Symanzik improvement, methods have been developed to remove  $\mathcal{O}(a)$  terms non-perturbatively using auxiliary simulations to tune parameters [10]. Such “non-perturbative improvement” is of great practical importance as it brings the discretization error from the fermion action down to the same level as that from the gauge action. It is used by essentially all simulations using clover fermions.

The advantages of Wilson fermions are their theoretical simplicity and relatively low computational cost. Their main disadvantage is the lack of chiral symmetry, which makes them difficult to use in cases where mixing with wrong chirality operators can occur, particularly if this involves divergences proportional to powers of  $1/a$ . A related problem is the presence of potential numerical instabilities due to spurious near-zero modes of the lattice Dirac operator. Ongoing work has, however, been successful at ameliorating these problems and increasing the range of quantities for which Wilson fermions can be used (see, e.g., Refs. 11–13).

Twisted-mass fermions [14] are a variant of Wilson fermions in which two flavors are treated together with an isospin-breaking mass term (the “twisted mass” term). The main advantage of this approach is that all errors linear in  $a$  are automatically removed (without the need for tuning of parameters) by a clever choice of twisted mass and operators [15]. A disadvantage is the presence of isospin breaking effects (such as a splitting between charged and neutral pion masses even when up and down quarks are degenerate), which, however, vanish as  $a^2 \Lambda^2$  in the continuum limit. Strange and charm quarks can be added as a second pair, with a term added to split their masses.

Staggered fermions are a reduced version of naive fermions in which there is only a single fermion Dirac component on each lattice site, with the full Dirac structure built up from neighboring sites [16]. They have the advantages of being somewhat faster to simulate than Wilson-like fermions, of preserving some chiral symmetry, and of having discretization errors of  $\mathcal{O}(a^2)$ . Their disadvantage is that they retain some of the doublers (3 for  $d = 4$ ). The action thus describes four degenerate fermions in the continuum limit. These are usually called “tastes”, to distinguish them from physical flavors, and the corresponding SU(4) symmetry is referred to as the “taste symmetry”. The preserved chiral symmetry in this formulation has non-singlet taste. Practical applications usually introduce one staggered fermion for each physical flavor, and remove contributions from the unwanted tastes by taking the fourth-root of the fermion determinant appearing in the path integral. The validity of this “rooting” procedure is not obvious because taste symmetry is violated for non-zero lattice spacing. Theoretical arguments, supported by numerical evidence, suggest that the procedure is valid as long as one takes the continuum limit before approaching the light quark mass region [17]. Additional issues arise for the valence quarks (those appearing in quark propagators, as described in Sec. 18.2 below), where rooting is not possible, and one must remove the extra tastes by hand [18].

Just as for Wilson fermions, the staggered action can be improved, so as to reduce discretization errors. The Asqtad ( $a$ -squared tadpole improved) action [19] was used until recently in many large scale simulations [20]. Most recent calculations use the HISQ (highly improved staggered quark) action, introduced in Ref. 21. This removes tree-level  $\mathcal{O}(a^2)$  errors, and leads to a substantial reduction in the breaking of taste symmetry, as well as a general reduction in the size of other discretization errors. It is tuned to reduce discretization errors for both light and heavier quarks, and is being used to directly simulate charm quarks.

There is an important class of lattice fermions, “Ginsparg-Wilsons fermions”, that possess a continuum-like chiral symmetry without introducing unwanted doublers. The lattice Dirac operator  $D$  for these fermions satisfies the Ginsparg-Wilson relation  $D\gamma_5 + \gamma_5 D = aD\gamma_5 D$  [22]. In the continuum, the right-hand-side vanishes, leading to chiral symmetry. On the lattice, it is non-vanishing, but with a particular form (with two factors of  $D$ ) that restricts the violations of chiral symmetry in Ward-Takahashi identities to short-distance terms that do not contribute to physical matrix elements [23]. In fact, one can define a modified chiral transformation on the lattice (by including dependence on the gauge fields) such that Ginsparg-Wilson fermions have an exact chiral symmetry for on-shell quantities [24]. The net result is that such fermions essentially have the same properties under chiral transformations as do continuum fermions, including the index theorem [23]. Their leading discretization errors are of  $\mathcal{O}(a^2)$ .

Two types of Ginsparg-Wilson fermions are currently being used in large-scale numerical simulations. The first is Domain-wall fermions (DWF). These are defined on a five-dimensional space, in which the fifth dimension is fictitious [25]. The action is chosen so that the



low-lying modes are chiral, with left- and right-handed modes localized on opposite four-dimensional surfaces. For an infinite fifth dimension, these fermions satisfy the Ginsparg-Wilson relation. In practice, the fifth dimension is kept finite, and there remains a small, controllable violation of chiral symmetry. The second type is Overlap fermions. These appeared from a completely different context and have an explicit form that exactly satisfies the Ginsparg-Wilson relation [26]. Their numerical implementation requires an approximation of the matrix sign function of a Wilson-like fermion operator, and various approaches are being used. In fact, it is possible to rewrite these approximations in terms of a five-dimensional formulation, showing that the DWF and Overlap approaches are essentially equivalent [27]. Numerically, the five-dimensional approach appears to be more computationally efficient.

The various lattice fermion formulations are often combined with the technique of link smearing. Here one couples the fermions to a smoother gauge link, defined by averaging with adjacent links in a gauge invariant manner. Several closely related implementations are being used. All reduce the coupling of fermions to the short-distance fluctuations in the gauge field, leading to an improvement in the numerical stability and speed of algorithms. One cannot perform this smearing too aggressively, however, since the smearing may distort short distance physics and enhance discretization errors.

As noted above, each fermion formulation has its own advantages and disadvantages. For instance, domain-wall and overlap fermions are theoretically preferred as they have chiral symmetry without doublers, but their computational cost is greater than for other choices. If the physics application of interest and the target precision do not require near-exact chiral symmetry, there is no strong motivation to use these expensive formulations. On the other hand, there is a class of applications (including the calculation of the  $\Delta I = 1/2$  amplitude for  $K \rightarrow \pi\pi$  decays and the S-parameter [28]) where chiral symmetry plays an essential role and for which the use of Ginsparg-Wilson fermions is strongly favored.

### 18.1.3. Heavy quarks on the lattice :

The fermion formulations described in the previous subsection can be used straightforwardly only for quarks whose masses are small compared to the lattice cutoff,  $m_q \lesssim 1/a$ . This is because there are discretization errors proportional to powers of  $am_q$ , and if  $am_q \gtrsim 1$  these errors are large and uncontrolled. Present LQCD simulations typically have cutoffs in the range of  $1/a = 2 - 4$  GeV (corresponding to  $a \approx 0.1 - 0.05$  fm). Thus, while for the up, down and strange quarks one has  $am_q \ll 1$ , for bottom quarks (with  $m_b \approx 4.5$  GeV) one must use alternative approaches. Charm quarks ( $m_c \approx 1.5$  GeV) are an intermediate case, allowing simulations using both direct and alternative approaches.

For the charm quark, the straightforward approach is to simultaneously reduce the lattice spacing and to improve the fermion action so as to reduce the size of errors proportional to powers of  $am_c$ . This approach has, for example, been followed successfully using the HISQ and twisted-mass actions [21,29,30]. It is important to note, however, that reducing  $a$  increases the computational cost because an increased number of lattice points are needed for the same physical volume. One cannot reduce the spatial size below  $2 - 3$  fm without introducing finite volume errors. Present lattices have typical sizes of  $\sim 64^3 \times 128$  (with the long direction being Euclidean time), and thus allow a lattice cutoff up to  $1/a \sim 4$  GeV.

Alternative approaches for discretizing heavy quarks are motivated by effective field theories. For a bottom quark in heavy-light hadrons, one can use Heavy Quark Effective Theory (HQET) to expand about the infinite quark-mass limit. In this limit, the bottom quark is a static color source, and one can straightforwardly write the corresponding lattice action [31]. Corrections, proportional to powers of  $1/m_b$ , can be introduced as operator insertions, with coefficients that can be determined non-perturbatively using existing techniques [32]. This method allows the continuum limit to be taken controlling all  $1/m_b$  corrections.

Another way of introducing the  $1/m_b$  corrections is to include the relevant terms in the effective action. This leads to a non-relativistic QCD (NRQCD) action, in which the heavy quark is described by a

two-component spinor [33]. This approach has the advantage over HQET that it can also be used for heavy-heavy systems, such as the Upsilon states. A disadvantage is that some of the parameters in this effective theory are determined perturbatively (originally at tree-level, but more recently at one-loop), which limits the precision of the final results. Although discretization effects can be controlled with good numerical precision for a range of lattice spacings, at fine enough lattice spacing the NRQCD effective theory no longer applies since power divergent terms become important, and taking the continuum limit would require fine-tuning a large number of couplings non-perturbatively.

This problem can be avoided if one uses HQET power counting to analyze and reduce discretization effects for heavy quarks while using conventional fermion actions [34]. For instance, one can tune the parameters of an improved Wilson quark action so that the leading HQET corrections to the static quark limit are correctly accounted for. As the lattice spacing becomes finer, the action smoothly goes over to that of a light Wilson quark action, where the continuum limit can be taken as usual. In principle, one can improve the action in the heavy quark regime up to arbitrarily high orders using HQET, but so far large-scale simulations have typically used clover improved Wilson quarks, where tuning the parameters of the action corresponds to including all corrections through next-to-leading order in HQET. Three different methods for tuning the parameters of the clover action are being used: the Fermilab [34], Tsukuba [35] and Columbia [36] approaches. An advantage of this HQET approach is that the  $c$  and  $b$  quarks can be treated on the same footing. Parameter tuning has typically been done perturbatively, as in NRQCD, but recent work using the Columbia approach has used non-perturbative tuning of some of the parameters [37,38].

Another approach is the “ratio method” introduced in [39]. Here one uses quarks with masses lying at, or slightly above, the charm mass  $m_c$ , which can be simulated with a relativistic action, and extrapolates to  $m_b$  incorporating the behavior predicted by HQET. The particular implementation relies on the use of ratios. As an example, consider the  $B$  meson decay constant  $f_B$ . According to HQET, this scales as  $1/\sqrt{m_B}$  for  $m_B \gg \Lambda_{\text{QCD}}$ , up to a logarithmic dependence that is calculable in perturbative QCD (but will be suppressed in the following). Here  $m_B$  is the  $B$  meson mass, which differs from  $m_b$  by  $\sim \Lambda_{\text{QCD}}$ . One considers the ratio  $y(\lambda, m_{b'}) \equiv f_{B''} \sqrt{m_{B''}} / f_{B'} \sqrt{m_{B'}}$  for fictitious  $B$  mesons containing  $b$  quarks with unphysical masses  $m_{b'}$  and  $m_{b''} = \lambda m_{b'}$ . HQET implies that  $y(\lambda, m_{b'})$  approaches unity for large  $m_{b'}$  and any fixed  $\lambda > 1$ . The ratios are evaluated on the lattice for the sequence of masses  $m_{b'} = m_c, \lambda m_c, \lambda^2 m_c$ , all well below the physical  $m_b$ , and for each the continuum limit is taken. The form of the ratio for larger values of  $m_{b'}$  is obtained by fitting, incorporating the constraints implied by HQET. The result for  $f_B \sqrt{m_B}$  is then obtained as a product of  $y$ 's with  $f_D \sqrt{m_D}$ .

### 18.1.4. QED on the lattice :

Quarks in nature are electrically charged, and the resultant coupling to photons leads to shifts in the properties of hadrons that are generically of  $\mathcal{O}(\alpha_{\text{EM}})$ . Thus, for example, the proton mass is increased by  $\sim 1$  MeV relative to that of the neutron due to its overall charge although this effect is more than compensated for by the  $\sim 2.5$  MeV relative decrease due to the up quark being lighter than the down quark [40]. This example shows that once pure QCD, isospin-symmetric lattice calculations reach percent level accuracy, further improvement requires the inclusion of effects due to both electromagnetism and the up-down mass difference. This level of accuracy has in fact been obtained for various quantities, e.g. light hadron masses and decay constants (see Ref. 41), and simulations including QED in addition to QCD are now beginning.

The extension of lattice methods to include QED is straightforward, although some new subtleties arise. The essential change is that the quark must now propagate through a background field containing both gluons and photons. The gauge field  $U_\mu$  that appears in covariant derivative Eq. (18.3) is extended from an SU(3) matrix to one living in U(3):  $U_\mu \rightarrow U_\mu e^{iaq_e A_\mu^{\text{EM}}}$ . Here  $A_\mu^{\text{EM}}$  is the photon field,  $e$  the electromagnetic coupling, and  $q$  the charge of the quark, e.g.  $q = 2/3$  for up and  $-1/3$  for down and strange quarks. The lattice action



for the photon that is typically used is a discretized version of the continuum action Eq. (18.2), rather than the form used for the gluons, Eq. (18.1). This “non-compact” action has the advantage that it is quadratic in  $A_\mu^{\text{EM}}$ , which simplifies the QED part of the generation of configurations.

One subtlety that arises is that Gauss’ law forbids a charged particle in a box with periodic boundary conditions. This finite volume effect can be overcome by including a uniform background charge, and this can be shown to be equivalent to removing the zero-momentum mode from the photon field. This is an example of the enhanced finite-volume effects that arise in the presence of the massless photon.

Simulations including QED have progressed over the last few years, and now a full inclusion of QED has been achieved with almost physical quark masses [40]. Alternative approaches have also been considered: reweighting the QCD fields *a posteriori* [42,43], and keeping only the linear term in an expansion in  $\alpha_{\text{EM}}$  about the QCD only case [44]. In addition, some calculations have included QED effects for the valence quarks but not the sea quarks (the “electroquenched approximation”).

#### 18.1.5. Basic inputs for lattice calculations :

Since LQCD is nothing but a regularization of QCD, the renormalizability of QCD implies that the number of input parameters in LQCD is the same as for continuum QCD—the strong coupling constant  $\alpha_s = g^2/(4\pi)$ , the quark masses for each flavor, and the CP violating phase  $\theta$ . The  $\theta$  parameter is usually assumed to be zero, while the other parameters must be determined using experimental inputs.

**18.1.5.1. Lattice spacing:** In QCD, the coupling constant is a function of scale. With lattice regularization, this scale is the inverse lattice spacing  $1/a$ , and choosing the bare coupling constant is equivalent to fixing the lattice spacing.

In principle,  $a$  can be determined using any dimensionful quantity measured by experiments. For example, using the mass of hadron  $H$  one has  $a = (am_H)^{\text{lat}}/m_H^{\text{exp}}$ . (Of course, one must first tune the quark masses to their physical values, as discussed below.) In practice, one chooses quantities that can be calculated accurately on the lattice, and that are only weakly dependent on the light quark masses. The latter property minimizes errors from extrapolating or interpolating to the physical light quark masses or from mistuning of these masses. Commonly used choices are the spin-averaged 1S-1P or 1S-2S splittings in the Upsilon system, the mass of the  $\Omega^-$  baryon, and the pion decay constant  $f_\pi$ . Ultimately, all choices must give consistent results for  $a$ , and that this is the case provides a highly non-trivial check of both the calculational method and of QCD.

#### 18.1.5.2. Light quark masses:

In LQCD simulations, the up, down and strange quarks are usually referred to as the light quarks, in the sense that  $m_q < \Lambda_{\text{QCD}}$ . (The standard definition of  $\Lambda_{\text{QCD}}$  is given in the “Quantum Chromodynamics” review; in this review we are using it only to indicate the approximate non-perturbative scale of QCD.) This condition is stronger than that used above to distinguish quarks with small discretization errors,  $m_q < 1/a$ . Loop effects from light quarks must be included in the simulations to accurately represent QCD. At present, most simulations are done in the isospin symmetric limit  $m_u = m_d \equiv m_\ell < m_s$ , and are often referred to as “ $N_f = 2 + 1$ ” simulations. Increasingly, simulations also include loops of charm quarks (denoted  $N_f = 2 + 1 + 1$  simulations), although the effect of charmed sea quarks on low-energy physics is generically expected to be at the sub-percent level [45]. Precision is now reaching the point where isospin breaking effects must be included. To do so without approximation requires simulating with nondegenerate up and down quarks (leading to  $N_f = 1 + 1 + 1$  or  $1 + 1 + 1 + 1$  simulations) as well as including electromagnetism (as described above). This has been done in Ref. 40. Alternatively, one can use a perturbative approach, expanding about the isospin symmetric theory and working to linear order in  $\alpha_{\text{EM}}$  and  $m_u - m_d$  [44].

We now describe the tuning of  $m_\ell$ ,  $m_s$  and  $m_c$  to their physical values. (For brevity, we ignore isospin violation in the following

discussion.) The most commonly used quantities for these tunings are, respectively,  $m_\pi$ ,  $m_K$  and  $m_{D_s}$ . If the scale is being set by  $m_\Omega$ , then one adjusts the lattice quark masses until the ratios  $m_\pi/m_\Omega$ ,  $m_K/m_\Omega$  and  $m_{D_s}/m_\Omega$  take their physical values. In the past, most calculations needed to extrapolate to the physical value of  $m_\ell$  (typically using forms based on chiral perturbation theory [ChPT]), while simulating directly at or near to the physical values of  $m_s$  and  $m_c$ . Present calculations are increasingly done with physical or near physical values of  $m_\ell$ , requiring at most only a short extrapolation.

#### 18.1.5.3. Heavy quark masses:

The  $b$  quark is usually treated only as a valence quark, with no loop effects included. The errors introduced by this approximation can be estimated to be  $\sim \alpha_s(m_b)\Lambda_{\text{QCD}}^2/m_b^2$  and are likely to be very small. In the past, the same approximation has been made for the  $c$  quark, leading to errors  $\sim \alpha_s(m_c)\Lambda_{\text{QCD}}^2/m_c^2$ . (See [45] for a quantitative estimate of the effects of including the charm quark on some low energy physical quantities.) For high precision, however, dynamical charm quarks are necessary, and some of the most recent simulations now include them.

The  $b$  quark mass can be tuned by setting heavy-heavy ( $\Upsilon$ ) or heavy-light ( $B$ ) meson masses to their experimental values. Consistency between these two determinations provides an important check that the determination of parameters in the heavy quark lattice formulations is being done correctly (see, e.g. Ref. 46).

#### 18.1.6. Sources of systematic error :

Lattice results have statistical and systematic errors that must be quantified for any calculation in order for the result to be a useful input to phenomenology. The statistical error is due to the use of Monte Carlo importance sampling to evaluate the path integral (a method discussed below). There are, in addition, a number of systematic errors that are always present to some degree in lattice calculations, although the size of any given error depends on the particular quantity under consideration and the parameters of the ensembles being used. The most common lattice errors are reviewed below.

Although not strictly a systematic error, it is important to note that the presence of long autocorrelations in the sequence of lattice configurations generated by the Monte Carlo method can lead to underestimates of statistical errors [47]. It is known that the global topological charge of the gauge fields decorrelates very slowly with certain algorithms [48]. The effect of poorly sampling topological charge is expected to be most significant for the pion mass and related quantities [49,50]. This issue becomes more relevant as the precision of the final results increases.

**18.1.6.1. Continuum limit:** Physical results are obtained in the limit that the lattice spacing  $a$  goes to zero. The Symanzik effective theory determines the scaling of lattice artefacts with  $a$ . Most lattice calculations use improved actions with leading discretizations errors of  $\mathcal{O}(a^2\Lambda^2)$ ,  $\mathcal{O}(\alpha_s a^2\Lambda^2)$ , or  $\mathcal{O}(\alpha_s a\Lambda)$ , where  $\Lambda$  is a typical momentum scale in the system. Knowledge of the scaling of the leading discretization errors allows controlled extrapolation to  $a = 0$  when multiple lattice spacings are available, as in current state-of-the-art calculations. Residual errors arise from the exclusion of subleading  $a$  dependence from the fits.

For many quantities the typical momentum scale in the system is  $\sim \Lambda_{\text{QCD}} \approx 300$  MeV. Discretization errors are expected to be larger for quantities involving larger scales, for example form factors or decays involving particles with momenta larger than  $\Lambda_{\text{QCD}}$ .

**18.1.6.2. Infinite volume limit:** LQCD calculations are necessarily carried out in finite space-time boxes, leading to departures of physical quantities (masses, decay constants, etc.) from their measured, infinite volume values. These finite-volume shifts are an important systematic that must be estimated and minimized.

Typical lattices are asymmetric, with  $N_s$  points in the three spatial directions and  $N_t$  in the (Euclidean) temporal direction. The spatial and temporal sizes in physical units are thus  $L_s = aN_s$  and  $L_t = aN_t$ , respectively. (Anisotropic lattice spacings are also sometimes used, as

discussed below in Sec. 1.3.1.) Typically,  $L_t \geq 2L_s$ , a longer temporal direction being used to allow excited-state contributions to correlators to decay. This means that the dominant impact of using finite volume is from the presence of a finite spatial box.

High-precision LQCD calculations are of quantities involving no more than a single particle in initial and final states (with the exception of the  $K \rightarrow \pi\pi$  decay amplitudes). For such quantities, once the volume exceeds about 2 fm (so that the particle is not “squeezed”), the dominant finite-volume effect comes from virtual pions wrapping around the lattice in the spatial directions. This effect is exponentially suppressed as the volume becomes large, roughly as  $\sim \exp(-m_\pi L_s)$ , and has been estimated using ChPT [51] or other methods [52]. The estimates suggest that finite volume shifts are sub-percent effects when  $m_\pi L_s \gtrsim 4$ , and most large-scale simulations use lattices satisfying this condition. This becomes challenging as one approaches the physical pion mass, for which  $L_s \gtrsim 5$  fm is required. At present, this can only be achieved by using relatively coarse lattices,  $a \gtrsim 0.07$  fm.

Finite volume errors are usually determined by repeating the simulations on two or more different volumes (with other parameters fixed). If different volumes are not available, the ChPT estimate can be used, often inflated to account for the fact that the ChPT calculation is truncated at some order.

In the future, LQCD calculations involving more than a single hadron will become increasingly precise. Examples include the calculation of resonance parameters and the above-mentioned  $K \rightarrow \pi\pi$  amplitudes. Finite volume effects are much larger in these cases, with power-law terms (e.g.  $1/L_s^3$ ) in addition to exponential dependence. Indeed, as will be discussed in Sec. 1.2.4., one can use the volume dependence to indirectly extract infinite-volume quantities such as scattering lengths. Doing so, however, requires a set of lattice volumes satisfying  $m_\pi L_s \gtrsim 4$  and is thus more challenging than for single-particle quantities.

#### 18.1.6.3. Chiral extrapolation:

Until recently, an important source of systematic error in LQCD calculations was the need to extrapolate in  $m_u$  and  $m_d$  (or, equivalently, in  $m_\pi$ ). This extrapolation was usually done using functional forms based on ChPT, or with analytic functions, with the difference between different fits used as an estimate of the systematic error, which was often substantial. Increasingly, however, calculations work directly at, or very close to, the physical quark masses. This either removes entirely, or greatly reduces, the uncertainties in the extrapolation, such that this error is subdominant.

#### 18.1.6.4. Operator matching:

Many of the quantities that LQCD can precisely calculate involve hadronic matrix elements of operators from the electroweak Hamiltonian. Examples include the pion and kaon decay constants, semileptonic form factors and the kaon mixing parameter  $B_K$  (the latter defined in Eq. (18.13)). The operators in the lattice matrix elements are defined in the lattice regularization scheme. To be used in tests of the Standard Model, however, they must be matched to the continuum regularization scheme in which the corresponding Wilson coefficients have been calculated. The only case in which such matching is not needed is if the operator is a conserved or partially conserved current. Similar matching is also needed for the conversion of lattice bare quark masses to those in the continuum  $\overline{\text{MS}}$  scheme.

Three methods are used to calculate the matching factors: perturbation theory (usually to one- or two-loop order), non-perturbative renormalization (NPR) using Landau-gauge quark and gluon propagators [53], and NPR using gauge-invariant methods based on the Schrödinger functional [54]. The NPR methods replace truncation errors (which can only be approximately estimated) by statistical and systematic errors which can be determined reliably and systematically reduced.

A common issue that arises in many such calculations (e.g. for quark masses and  $B_K$ ) is that, using NPR, one ends up with operators regularized in a MOM-like (or Schrödinger functional) scheme, rather than the  $\overline{\text{MS}}$  scheme mostly used for calculating the Wilson coefficients. To make contact with this scheme requires a purely continuum perturbative matching calculation. The resultant

truncation error can, however, be minimized by pushing up the momentum scale at which the matching is done using step-scaling techniques as part of the NPR calculation [55]. It should also be noted that this final step in the conversion to the  $\overline{\text{MS}}$  scheme could be avoided if continuum calculations used a MOM-like scheme.

## 18.2. Methods and status

Once the lattice action is chosen, it is straightforward to define the quantum theory using the path integral formulation. The Euclidean-space partition function is

$$Z = \int [dU] \prod_f [dq_f][d\bar{q}_f] e^{-S_g[U] - \sum_f \bar{q}_f (D[U] + m_f) q_f}, \quad (18.4)$$

where link variables are integrated over the SU(3) manifold,  $q_f$  and  $\bar{q}_f$  are Grassmann (anticommuting) quark and antiquark fields of flavor  $f$ , and  $D[U]$  is the chosen lattice Dirac operator with  $m_f$  the quark mass in lattice units. Integrating out the quark and antiquark fields, one arrives at a form suitable for simulation:

$$Z = \int [dU] e^{-S_g[U]} \prod_f \det(D[U] + m_f). \quad (18.5)$$

The building blocks for calculations are expectation values of multi-local gauge-invariant operators, also known as “correlation functions”,

$$\langle \mathcal{O}(U, q, \bar{q}) \rangle = (1/Z) \int [dU] \prod_f [dq_f][d\bar{q}_f] \mathcal{O}(U, q, \bar{q}) e^{-S_g[U] - \sum_f \bar{q}_f (D[U] + m_f) q_f}. \quad (18.6)$$

If the operators depend on the (anti-)quark fields  $q_f$  and  $\bar{q}_f$ , then integrating these fields out leads not only to the fermion determinant but also, through Wick’s theorem, to a series of quark “propagators”,  $(D[U] + m_f)^{-1}$ , connecting the positions of the fields.

This set-up allows one to choose, by hand, the masses of the quarks in the determinant (the sea quarks) differently from those in the propagators (valence quarks). This is called “partial quenching”, and is used by some calculations as a way of obtaining more data points from which to extrapolate both sea and valence quarks to their physical values.

#### 18.2.1. Monte-Carlo method:

Since the number of integration variables  $U$  is huge ( $N_s^3 \times N_t \times 4 \times 9$ ), direct numerical integration is impractical and one has to use Monte-Carlo techniques. In this method, one generates a Markov chain of gauge configurations (a “configuration” being the set of  $U$ ’s on all links) distributed according to the probability measure  $[dU] e^{-S_g[U]} \prod_f \det(D[U] + m_f)$ . Once the configurations are generated, expectation values  $\langle \mathcal{O}(U, q, \bar{q}) \rangle$  are calculated by averaging over those configurations. In this way the configurations can be used repeatedly for many different calculations, and there are several large collections of ensembles of configurations (with a range of values of  $a$ , lattice sizes and quark masses) that are publicly available through the International Lattice Data Grid (ILDG). As the number of the configurations,  $N$ , is increased, the error decreases as  $1/\sqrt{N}$ .

The most challenging part of the generation of gauge configurations is the need to include the fermion determinant. Direct evaluation of the determinant is not feasible, as it requires  $\mathcal{O}((N_s^3 \times N_t)^3)$  computations. Instead, one rewrites it in terms of “pseudofermion” fields  $\phi$  (auxiliary fermion fields with bosonic statistics). For example, for two degenerate quarks one has

$$\det(D[U] + m_f)^2 = \int [d\phi] e^{-\phi^\dagger (D[U] + m_f)^{-2} \phi}. \quad (18.7)$$

By treating the pseudofermions as additional integration variables in the path integral, one obtains a totally bosonic representation. The price one pays is that the pseudofermion effective action is highly non-local since it includes the inverse Dirac operator  $(D[U] + m_f)^{-1}$ .

Thus, the large sparse matrix  $(D[U] + m)$  has to be inverted every time one needs an evaluation of the effective action.

Present simulations generate gauge configurations using the Hybrid Monte Carlo (HMC) algorithm [56], or variants thereof. This algorithm combines molecular dynamics (MD) evolution in a fictitious time (which is also discretized) with a Metropolis “accept-reject” step. It makes a global update of the configuration, and is made exact by the Metropolis step. In its original form it can be used only for two degenerate flavors, but extensions (particularly the rational HMC [57]) are available for single flavors. Considerable speed-up of the algorithms has been achieved over the last two decades using a variety of techniques.

All these algorithms spend the bulk of their computational time on the repeated inversion of  $(D[U] + m)$  acting on a source (which is required at every step of the MD evolution). Inversions are done using a variety of iterative algorithms, *e.g.* the conjugate gradient algorithm. In this class of algorithms, computational cost is proportional to the condition number of the matrix, which is the ratio of maximum and minimum eigenvalues. For  $(D[U] + m)$  the smallest eigenvalue is  $\approx m$ , so the condition number and cost are inversely proportional to the quark mass. This is a major reason why simulations at the physical quark mass are challenging. Recent algorithmic studies are making progress in significantly reducing this problem.

A practical concern is the inevitable presence of correlations between configurations in the Markov chain. These are characterized by an autocorrelation length in the fictitious MD time. One aims to use configurations separated in MD time by greater than this autocorrelation length. In practice, it is difficult to measure this length accurately, and this leads to some uncertainty in the resulting statistical errors, as well as the possibility of insufficient equilibration.

For most of the applications of LQCD discussed in this review, the cost of generating gauge configurations is larger than or similar to that of performing the “measurements” on those configurations. The computational cost of gauge generation grows with the lattice volume,  $V_{\text{lat}} = N_s^3 N_t$ , as  $V_{\text{lat}}^{1+\delta}$ . Here  $\delta = 1/4$  for the HMC algorithm [58] and can be reduced slightly using modern variants. Such growth with  $V_{\text{lat}}$  provides a (time-dependent) limit on the largest lattice volumes that can be simulated. At present, the largest lattices being used have  $N_s = 144$  and  $N_t = 288$ . Typically one aims to create an ensemble of  $\sim 10^3$  statistically independent configurations at each choice of parameters ( $a$ ,  $m_q$  and  $V_{\text{lat}}$ ). For most physical quantities of interest, this is sufficient to make the resulting statistical errors smaller than or comparable to the systematic errors.

### 18.2.2. Two-point functions :

One can extract properties of stable hadrons using two-point correlation functions,  $\langle O_X(x) O_Y^\dagger(0) \rangle$ . Here  $O_{X,Y}(x)$  are operators that have non-zero overlaps with the hadronic state of interest  $|H\rangle$ , *i.e.*  $\langle 0|O_{X,Y}(x)|H\rangle \neq 0$ . One usually Fourier-transforms in the spatial directions and considers correlators as a function of Euclidean time:

$$C_{XY}(t; \vec{p}) = \sum_{\vec{x}} \langle O_X(t, \vec{x}) O_Y^\dagger(0) \rangle e^{-i\vec{p} \cdot \vec{x}}. \quad (18.8)$$

(Here and throughout this section all quantities are expressed in dimensionless lattice units, so that, for example,  $\vec{p} = a\vec{p}_{\text{phys}}$ .) By inserting a complete set of states having spatial momentum  $\vec{p}$ , the two-point function can be written as

$$C_{XY}(t; \vec{p}) = \sum_{i=0}^{\infty} \frac{1}{2E_i(\vec{p})} \langle 0|O_X(0)|H_i(\vec{p})\rangle \langle H_i(\vec{p})|O_Y^\dagger(0)|0\rangle e^{-E_i(\vec{p})t}, \quad (18.9)$$

where the energy of the  $i$ -th state  $E_i(\vec{p})$  appears as an eigenvalue of the time evolution operator  $e^{-Ht}$  in the Euclidean time direction. The factor of  $1/[2E_i(\vec{p})]$  is due to the relativistic normalization used for the states. For large enough  $t$ , the dominant contribution is that of the lowest energy state  $|H_0(\vec{p})\rangle$ :

$$C_{XY}(t) \xrightarrow{t \rightarrow \infty} \frac{1}{2E_0(\vec{p})} \langle 0|O_X(0)|H_0(\vec{p})\rangle \langle H_0(\vec{p})|O_Y^\dagger(0)|0\rangle e^{-E_0(\vec{p})t}. \quad (18.10)$$

One can thus obtain the energy  $E_0(\vec{p})$ , which equals the hadron mass  $m_H$  when  $\vec{p} = 0$ , and the product of matrix elements  $\langle 0|O_X(0)|H_i(\vec{p})\rangle \langle H_i(\vec{p})|O_Y^\dagger(0)|0\rangle$ .

This method can be used to determine the masses of all the stable mesons and baryons by making appropriate choices of operators. For example, if one uses the axial current,  $O_X = O_Y = A_\mu = \bar{d}\gamma_\mu\gamma_5 u$ , then one can determine  $m_{\pi^+}$  from the rate of exponential fall-off, and in addition the decay constant  $f_\pi$  from the coefficient of the exponential. A complication arises for states with high spins ( $j \geq 4$  for bosons) because the spatial rotation group on the lattice is a discrete subgroup of the continuum group  $\text{SO}(3)$ . This implies that lattice operators, even when chosen to lie in irreducible representations of the lattice rotation group, have overlap with states that have a number of values of  $j$  in the continuum limit [59]. For example  $j = 0$  operators can also create mesons with  $j = 4$ . A method to overcome this problem has recently been introduced [60,61].

The expression given above for the correlator  $C_{XY}(t; \vec{p})$  shows how, in principle, one can determine the energies of the excited hadron states having the same quantum numbers as the operators  $O_{X,Y}$ , by fitting the correlation function to a sum of exponentials. In practice, this usually requires using a large basis of operators and adopting the variational approach such as that of Ref. 62. One can also use an anisotropic lattice in which  $a_t$ , the lattice spacing in the time direction, is smaller than its spatial counterpart  $a_s$ . This allows better separation of the different exponentials. Using a combination of these and other technical improvements extensive excited-state spectra have recently been obtained [61,63,64].

### 18.2.3. Three-point functions :

Hadronic matrix elements needed to calculate semileptonic form factors and neutral meson mixing amplitudes can be computed from three-point correlation functions. We discuss here, as a representative example, the  $D \rightarrow K$  amplitude. As in the case of two-point correlation functions one constructs operators  $O_D$  and  $O_K$  having overlap, respectively, with the  $D$  and  $K$  mesons. We are interested in calculating the matrix element  $\langle K|V_\mu|D\rangle$ , with  $V_\mu = \bar{c}\gamma_\mu s$  the vector current. To obtain this, we use the three-point correlator

$$C_{KV_\mu D}(t_x, t_y; \vec{p}) = \sum_{\vec{x}, \vec{y}} \langle O_K(t_x, \vec{x}) V_\mu(0) O_D^\dagger(t_y, \vec{y}) \rangle e^{-i\vec{p} \cdot \vec{x}}, \quad (18.11)$$

and focus on the limit  $t_x \rightarrow \infty$ ,  $t_y \rightarrow -\infty$ . In this example we set the  $D$ -meson at rest while the kaon carries three-momentum  $\vec{p}$ . Momentum conservation then implies that the weak operator  $V_\mu$  inserts three-momentum  $-\vec{p}$ . Inserting a pair of complete sets of states between each pair of operators, we find

$$C_{KV_\mu D}(t_x, t_y; \vec{p}) = \sum_{i,j} \frac{1}{2m_{D_i} 2E_{K_j}(\vec{p})} e^{-m_{D_i} t_x - E_{K_j}(\vec{p}) |t_y|} \times \\ \times \langle 0|O_K(t_x, \vec{x})|K_i(\vec{p})\rangle \langle K_i(\vec{p})|V_\mu(0)|D_j(\vec{0})\rangle \langle D_j(\vec{0})|O_D^\dagger(0)|0\rangle. \quad (18.12)$$

The matrix element  $\langle K_i(\vec{p})|V_\mu(0)|D_j(\vec{0})\rangle$  can then be extracted, since all other quantities in this expression can be obtained from two-point correlation functions. Typically one is interested in the weak matrix elements of ground states, such as the lightest pseudoscalar mesons. In the limit of large separation between the three operators in Euclidean time, the three-point correlation function yields the weak matrix element of the transition between ground states.

### 18.2.4. Scattering amplitudes and resonances :

The methods described thus far yield matrix elements involving single, stable particles (where by stable here we mean absolutely stable to strong interaction decays). Most of the particles listed in the Review of Particle Properties are, however, unstable—they are resonances decaying into final states consisting of multiple strongly interacting particles. LQCD simulations cannot directly calculate resonance properties, but methods have been developed to do so indirectly for resonances coupled to two-particle final states in the elastic regime [65].

The difficulty faced by LQCD calculations is that, to obtain resonance properties, or, more generally, scattering phase-shifts, one

must calculate multiparticle scattering amplitudes in momentum space and put the external particles on their mass-shells. This requires analytically continuing from Euclidean to Minkowski momenta. Although it is straightforward in LQCD to generalize the methods described above to calculate four- and higher-point correlation functions, one necessarily obtains them at a discrete and finite set of Euclidean momenta. Analytic continuation to  $p_E^2 = -m^2$  is then an ill-posed and numerically unstable problem. The same problem arises for single-particle states, but can be largely overcome by picking out the exponential fall-off of the Euclidean correlator, as described above. With a multi-particle state, however, there is no corresponding trick, except for two particles at threshold [66].

What LQCD can calculate are the energies of the eigenstates of the QCD Hamiltonian in a finite box. The energies of states containing two stable particles, e.g. two pions, clearly depend on the interactions between the particles. It is possible to invert this dependence and, with plausible assumptions, determine the scattering phase-shifts at a discrete set of momenta from a calculation of the two-particle energy levels for a variety of spatial volumes [65]. This is a challenging calculation, but it has recently been carried through in several channels with quark masses approaching physical values. Channels studied include  $\pi\pi$  (for  $I = 2, 1$  and  $0$ ),  $K\pi$ ,  $KD$  and  $DD^*$ . For recent reviews see Ref. 67. Extensions to nucleon interactions are also being actively studied [68]. The generalization of the formalism to the case of three particles is under active consideration [69].

It is also possible to extend the methodology to calculate electroweak decay amplitudes to two particles below the inelastic threshold, e.g.  $\Gamma(K \rightarrow \pi\pi)$  [70]. Results for both the  $\Delta I = 3/2$  and  $1/2$  amplitudes with physical quark masses have been obtained [71], the former now including a controlled continuum limit [72]. Partial extensions of the formalism above the elastic threshold have been worked out, in particular for the case of multiple two-particle channels [73]. An extension to decays with many multiparticle channels, e.g. hadronic  $B$  decays, has, however, yet to be formulated.

**18.2.5. Recent advances :** In some physics applications, one is interested in the two-point correlation function  $\langle O_X(x) O_Y^\dagger(0) \rangle$  for all values of the separation  $x$ , not just its asymptotic form for large separations (which is used to determine the hadron spectrum as sketched above). A topical example is the hadronic vacuum polarization function  $\Pi_{\mu\nu}(x) = \langle V_\mu(x) V_\nu(0) \rangle$  and its Fourier transform  $\Pi_{\mu\nu}(q^2)$ . Since the lattice is in Euclidean space-time, only space-like momenta,  $q^2 = -Q^2 < 0$ , are accessible. Nevertheless, this quantity is of significant interest. It is related by a dispersion relation to the cross section for  $e^+ + e^- \rightarrow$  hadrons, and is needed for a first-principles calculation of the “hadronic vacuum polarization” contribution to the muon anomalous magnetic moment  $a_\mu$ . This is the contribution with the largest theoretical uncertainty at present. There are a number of lattice calculations of this contribution (see, e.g., [74–77] following the pioneering work of Ref. 78).

Exploratory calculations of the light-by-light scattering contribution to  $a_\mu$  are also underway. These involve the evaluation of a four-point correlation function (see, e.g., Ref. 79). Another process under consideration is the long-distance contribution to the neutral kaon mass splitting,  $\Delta M_K$ . This also requires the evaluation of a four-point function, constructed from the two-point functions described above by the insertion of two electroweak Hamiltonians [80].

#### 18.2.6. Status of LQCD simulations :

Until the 1990s, most large-scale lattice simulations were limited to the “quenched” approximation, wherein the fermion determinant is omitted from the path integral. While much of the basic methodology was developed in this era, the results obtained had uncontrolled systematic errors and were not suitable for use in placing precision constraints on the Standard Model. During the 1990s, more extensive simulations including the fermion determinant (also known as simulations with “dynamical” fermions) were begun, but with unphysically heavy quark masses ( $m_\ell \sim 50 - 100$  MeV), such that the extrapolation to the physical light quark masses was a source of large systematic errors [81]. During the 2000s, advances in both algorithms and computers allowed simulations to reach much smaller

quark masses ( $m_\ell \sim 10 - 20$  MeV) such that LQCD calculations of selected quantities with all sources of error controlled and small became available. Their results played an important role in constraints on the CKM matrix and other phenomenological analyses. In the last few years, simulations directly at the physical isospin-symmetric light quark masses have become standard, removing the need for a chiral extrapolation and thus significantly reducing the overall error. The present frontier, as noted above, is the inclusion of isospin breaking. This will be needed to push the accuracy of calculations below the percent level.

On a more qualitative level, analytic and numerical results from LQCD have demonstrated that QCD confines color and spontaneously breaks chiral symmetry. Confinement can be seen as a linearly rising potential between heavy quark and anti-quark in the absence of quark loops. Analytically, this can be shown in the strong coupling limit  $g_{\text{lat}} \rightarrow \infty$  [1]. At weaker couplings there are precise numerical calculations of the potential that clearly show that this behavior persists in the continuum limit [82,83,84].

Chiral symmetry breaking was also demonstrated in the strong coupling limit on the lattice [16,85], and there have been a number of numerical studies showing that this holds also in the continuum limit. The accumulation of low-lying modes of the Dirac operator, which is the analog of Cooper pair condensation in superconductors, has been observed, yielding a determination of the chiral condensate [86–90]. Many relations among physical quantities that can be derived under the assumption of broken chiral symmetry have been confirmed by a number of lattice groups [41].

### 18.3. Physics applications

In this section we describe the main applications of LQCD that are both computationally mature and relevant for the determination of particle properties.

A general feature to keep in mind is that, since there are many different choices for lattice actions, all of which lead to the same continuum theory, a crucial test is that results for any given quantity are consistent. In many cases, different lattice calculations are completely independent and often have very different systematic errors. Thus final agreement, if found, is a highly non-trivial check, just as it is for different experimental measurements.

The number, variety and precision of the calculations has progressed to the point that an international “Flavour Lattice Averaging Group” (FLAG) has been formed. The main aims of FLAG include collecting all lattice results of relevance for a variety of phenomenologically interesting quantities and providing averages of those results which pass appropriate quality criteria. The averages attempt to account for possible correlations between results (which can arise, for example, if they use common gauge configurations). The quantities considered are those we discuss in this section, with the exception of the hadron spectrum. The most recent FLAG review is from 2013 [41] (with an update due in 2015-16). The interested reader can consult the FLAG review for very extensive discussions of the details of the calculations and of the sources of systematic errors.

We stress that the results we quote below are those obtained using the physical complement of light quarks (i.e.  $N_f = 2 + 1$  or  $2 + 1 + 1$  simulations).

#### 18.3.1. Spectrum :

The most basic prediction of LQCD is of the hadron spectrum. Once the input parameters are fixed as described in Sec. 18.1.5, the masses or resonance parameters of all other states can be predicted. This includes hadrons composed of light ( $u$ ,  $d$  and  $s$ ) quarks, as well as heavy-light and heavy-heavy hadrons. It also includes quark-model exotics (e.g.  $J^{PC} = 1^{-+}$  mesons) and glueballs. Thus, in principle, LQCD calculations should be able to reproduce many of the experimental results compiled in the Review of Particle Properties. Doing so would test both that the error budgets of LQCD calculations are accurate and that QCD indeed describes the strong interactions in the low-energy domain. The importance of the latter test can hardly be overstated.

What is the status of this fundamental test? As discussed in Sec. 1.2, LQCD calculations are most straightforward for stable, low-lying hadrons. Calculations of the properties of resonances which can decay into only two particles are more challenging, though substantial progress has been made. First theoretical work on decays to more than two particles has begun, but the methodology is not yet practical. It is also more technically challenging to calculate masses of flavor singlet states (which can annihilate into purely gluonic intermediate states) than those of flavor non-singlets, although again algorithmic and computational advances have begun to make such calculations accessible, although not yet for physical quark masses. The present status for light hadrons is that fully controlled results are available for the masses of the octet light baryons, while results with less than complete control are available for the decuplet baryon resonances, the vector meson resonances and the  $\eta$  and  $\eta'$ . In addition, it has been possible to calculate the isospin splitting in light mesons and baryons (due to the up-down mass difference and the incorporation of QED) [40]. There are also extensive results for heavy-light ( $D$  and  $B$  systems) and heavy-heavy ( $J/\psi$  and  $\Upsilon$  systems). All present results, which are discussed in the “Quark Model” review, are consistent with experimental values, and several predictions have been made. For a recent extensive review of lattice results see also Ref. 91.

### 18.3.2. Decay constants and bag parameters :

The pseudoscalar decay constants can be determined from two-point correlation functions involving the axial-vector current, as discussed in Sec. 18.2.2. The decay constant  $f_P$  of a meson  $P$  is extracted from the weak matrix element involving the axial-vector current using the relation  $\langle 0|A_\mu(x)|P(\vec{p})\rangle = f_P p_\mu \exp(-ip \cdot x)$ , where  $p_\mu$  is the momentum of  $P$  and  $A_\mu(x)$  is the axial-vector current. Since they are among the simplest quantities to calculate, decay constants provide good benchmarks for lattice methods, in addition to being important inputs for flavor physics phenomenology in their own right. Results from many lattice groups for the pion and kaon decay constants now have errors at the percent level or better. The decay constants in the charm and bottom sectors,  $f_D$ ,  $f_{D_s}$ ,  $f_B$ , and  $f_{B_s}$ , have also been calculated to high precision. Lattice results for all of these decay constants are discussed in detail in the review “Leptonic Decays of Charged Pseudoscalar Mesons.”

Another important lattice quantity is the kaon bag parameter  $B_K$ , which is needed to turn the precise measurement of CP-violation in kaon mixing into a constraint on the Standard Model. It is defined by

$$\frac{8}{3} m_K^2 f_K^2 B_K(\mu) = \langle \bar{K}^0 | Q_{\Delta S=2}(\mu) | K^0 \rangle, \quad (18.13)$$

where  $m_K$  is the kaon mass,  $f_K$  is the kaon decay constant,  $Q_{\Delta S=2} = \bar{s}\gamma_\mu(1-\gamma_5)d\bar{s}\gamma_\mu(1-\gamma_5)d$  is the four-quark operator of the effective electroweak Hamiltonian and  $\mu$  is the renormalization scale. The short distance contribution to the electroweak Hamiltonian can be calculated perturbatively, but the hadronic matrix element parameterized by  $B_K$  must be computed using non-perturbative methods. In order to be of use to phenomenology, the renormalization factor of the four-quark operator must be matched to a continuum renormalization scheme, e.g. to  $\overline{\text{MS}}$ , as described in Sec. 18.1.6.4. Determinations with percent-level precision using different fermion actions and  $N_f = 2 + 1$  light sea quarks are now available using DWF [92], staggered fermions [93], DWF valence on staggered sea quarks [94], and Wilson fermions [12]. The results are all consistent. Based on results available in 2013, FLAG quoted an average of  $\hat{B}_K = 0.766(10)$  [41]. The updates of Refs. 92 and 93 are consistent with this result.

The bag parameters for  $B$  and  $B_s$  meson mixing are defined analogously to that for kaon mixing. The  $B$  and  $B_s$  mesons contain a valence  $b$ -quark so that calculations of these quantities must use one of the methods for heavy quarks described above. Calculations with  $N_f = 2 + 1$  light fermions have been done using NRQCD [95], the Fermilab formalism [96], and static heavy quarks [97]. All results are consistent. The FLAG averages from 2013 for the quantities relevant for  $B_s$  and  $B$  mixing are  $f_{B_s}\sqrt{B_{B_s}} = 266(18)$  MeV and  $f_B\sqrt{B_B} = 216(15)$  MeV, with their ratio (which is somewhat better determined) being  $\xi = 1.268(63)$  [41]. Note that the errors for

quantities involving  $b$  quarks are larger than those for quantities involving only light quarks.

The results for mixing matrix elements are used in the reviews “The CKM Quark-Mixing Matrix,” and “ $B_0 - \bar{B}_0$  Mixing.”

### 18.3.3. Form factors ( $K \rightarrow \pi\ell\nu$ , $D \rightarrow K\ell\nu$ , $B \rightarrow \pi\ell\nu$ , $B \rightarrow D^{(*)}\ell\nu$ ) :

Semileptonic decay rates can be used to extract CKM matrix elements once the semileptonic form factors are known from lattice calculations. For example, the matrix element of a pseudoscalar meson  $P$  undergoing semileptonic decay to another pseudoscalar meson  $D$  is mediated by the vector current, and can be written in terms of form factors as

$$\langle D(p_D) | V_\mu | P(p_P) \rangle = f_+(q^2)(p_D + p_P - \Delta)_\mu + f_0(q^2)\Delta_\mu, \quad (18.14)$$

where  $q = p_D - p_P$ ,  $\Delta_\mu = (m_D^2 - m_P^2)q_\mu/q^2$  and  $V_\mu$  is the quark vector current. The shape of the form factor is typically well determined by experiment, and the value of  $f_+(q^2)$  at some reference value of  $q^2$  is needed from the lattice in order to extract CKM matrix elements. Typically  $f_+(q^2)$  dominates the decay rate, since the contribution from  $f_0(q^2)$  is suppressed when the final state lepton is light.

The form factor  $f_+(0)$  for  $K \rightarrow \pi\ell\nu$  decays is highly constrained by the Ademollo-Gatto theorem [98] and chiral symmetry. Old estimates using chiral perturbation theory combined with quark models quote sub-percent precision [99], though they suffer from some model dependence. Utilizing the constraint from the vector current conservation that  $f_+(0)$  is normalized to unity in the limit of degenerate up and strange quark masses, the lattice calculation can be made very precise and has now matched the precision of the phenomenological estimates [100–106]. The FLAG average from its 2013 edition is  $f_+(0) = 0.967(4)$  [41].

Charm meson semileptonic decays have been calculated by different groups using methods similar to those used for charm decay constants, and results are steadily improving in precision [107,108]. For semileptonic decays involving a bottom quark, one uses HQET or NRQCD to control the discretization errors of the bottom quark. The form factors for the semileptonic decay  $B \rightarrow \pi\ell\nu$  have been calculated in unquenched lattice QCD by a number of groups [109–111]. These  $B$  semileptonic form factors are difficult to calculate at low  $q^2$ , i.e. when the mass of the  $B$ -meson must be balanced by a large pion momentum, in order to transfer a small momentum to the lepton pair. The low  $q^2$  region has large discretization errors and very large statistical errors, while the high  $q^2$  region is much more accessible to the lattice. For experiment, the opposite is true. To combine lattice and experimental results it has proved helpful to use the  $z$ -parameter expansion [112]. This provides a theoretically constrained parameterization of the entire  $q^2$  range, and allows one to obtain  $|V_{ub}|$  without model dependence [113,114].

The semileptonic decays  $B \rightarrow D\ell\nu$  and  $B \rightarrow D^*\ell\nu$  can be used to extract  $|V_{cb}|$  once the corresponding form factors are known. At present only one unquenched calculation exists for the  $B \rightarrow D^*\ell\nu$  form factor, where the Fermilab formulation of the heavy quark was adopted [115,116]. This calculation is done at zero-recoil because that is where the lattice systematic errors are smallest. Calculations at non-zero recoil in unquenched lattice QCD have recently been done for the first time for the form factors needed to extract  $|V_{cb}|$  from  $B \rightarrow D\ell\nu$  decays [117].

The results discussed in this section are used in the reviews “The CKM Quark-Mixing Matrix,” “ $V_{ud}$ ,  $V_{us}$ , the Cabibbo Angle and CKM Unitarity,” and “ $V_{cb}$  and  $V_{ub}$  CKM Matrix Elements.”

### 18.3.4. Strong coupling constant :

As explained in Sec. 18.1.5.1, for a given lattice action, the choice of bare lattice coupling constant,  $g_{\text{lat}}$ , determines the lattice spacing  $a$ . If one then calculates  $a$  as described in Sec. 18.1.5.1, one knows the strong coupling constant in the bare lattice scheme at the scale  $1/a$ ,  $\alpha_{\text{lat}} = g_{\text{lat}}^2/(4\pi)$ . This is not, however, useful for comparing to results for  $\alpha_s$  obtained from other inputs, such as deep inelastic scattering or jet shape variables. This is because the latter results give  $\alpha_s$  in the  $\overline{\text{MS}}$  scheme, which is commonly used in such analyses, and the conversion

factor between these two schemes is known to converge extremely poorly in perturbation theory. Instead one must use a method which directly determines  $\alpha_s$  on the lattice in a scheme closer to  $\overline{\text{MS}}$ .

Several such methods have been used, all following a similar strategy. One calculates a short-distance quantity  $K$  both perturbatively ( $K^{\text{PT}}$ ) and non-perturbatively ( $K^{\text{NP}}$ ) on the lattice, and requires equality:  $K^{\text{NP}} = K^{\text{PT}} = \sum_{i=0}^n c_i \alpha_s^i$ . Solving this equation one obtains  $\alpha_s$  at a scale related to the quantity being used. Often,  $\alpha_s$  thus obtained is not defined in the conventional  $\overline{\text{MS}}$  scheme, and one has to convert among the different schemes using perturbation theory. Unlike for the bare lattice scheme, the required conversion factors are reasonably convergent. As a final step, one uses the renormalization group to run the resulting coupling to a canonical scale (such as  $M_Z$ ).

In the work of the HPQCD collaboration [118], the short-distance quantities are Wilson loops of several sizes and their ratios. These quantities are perturbatively calculated to  $\mathcal{O}(\alpha_s^3)$  using the  $V$ -scheme defined through the heavy quark potential. The coefficients of even higher orders are estimated using the data at various values of  $a$ .

Another choice of short-distance quantities is to use current-current correlators. Appropriate moments of these correlators are ultraviolet finite, and by matching lattice results to the *continuum* perturbative predictions, one can directly extract the  $\overline{\text{MS}}$  coupling. The JLQCD collaboration [119] uses this approach with light overlap fermions, while the HPQCD collaboration uses charm-quark correlators and HISQ fermions [120–122]. Yet another choice of short-distance quantity is the static-quark potential, where the lattice result for the potential is compared to perturbative calculations; this method was used to compute  $\alpha_s$  within 2+1 flavor QCD [123]. The ETM Collaboration obtains  $\alpha_s$  by a comparison of lattice data for the ghost-gluon coupling with that of perturbation theory [124], providing the first determination of  $\alpha_s$  with 2+1+1 flavors of dynamical quarks.

With a definition of  $\alpha_s$  given using the Schrödinger functional, one can non-perturbatively control the evolution of  $\alpha_s$  to high-energy scales, such as 100 GeV, where the perturbative expansion converges very well. This method developed by the ALPHA collaboration [55] has been applied to 2+1-flavor QCD in Refs. 125,126.

The various lattice methods for calculating  $\alpha_s$  have significantly different sources of systematic error. Thus the good agreement between the approaches (which can be seen in the “Quantum Chromodynamics” review) provides a strong check on the final result.

### 18.3.5. Quark masses :

Once the quark mass parameters are tuned in the lattice action, the remaining task is to convert them to those of the conventional definition. Since the quarks do not appear as asymptotic states due to confinement, the pole mass of the quark propagator is not a physical quantity. Instead, one defines the quark mass after subtracting the ultra-violet divergences in some particular way. The conventional choice is again the  $\overline{\text{MS}}$  scheme at a canonical scale such as 2 or 3 GeV. Ratios such as  $m_c/m_s$  and  $m_b/m_c$  are also useful as they are free from multiplicative renormalization (in a mass-independent scheme).

As discussed in Sec. 18.1.6.4, one must convert the lattice bare quark mass to that in the  $\overline{\text{MS}}$  scheme. Older calculations did so directly using perturbation theory; most recent calculations use an intermediate NPR method (e.g. RI/MOM or RI/SMOM) which is then converted to the  $\overline{\text{MS}}$  scheme using perturbation theory.

Alternatively, one can use a definition based on the Schrödinger functional, which allows one to evolve the quark mass to a high scale non-perturbatively [127]. In practice, one can reach scales as high as  $\sim 100$  GeV, at which matching to the  $\overline{\text{MS}}$  scheme can be reliably calculated in perturbation theory.

Another approach available for heavy quarks is to match current-current correlators at short distances calculated on the lattice to those obtained in continuum perturbation theory in the  $\overline{\text{MS}}$  scheme [120,121]. This has allowed an accurate determination of  $m_c$  and is also beginning to be used for  $m_b$  [121,122].

The ratio method for heavy quarks (discussed earlier) can also be used to determine  $m_b$  [128].

Results are summarized in the review of “Quark Masses”.

### 18.3.6. Other applications :

In this review we have concentrated on applications of LQCD that are relevant to the quantities discussed in the Review of Particle Properties. We have not discussed at all several other applications which are being actively pursued by simulations. Here we list the major such applications. The reader can consult the texts [2–4] for further details, as well as the proceedings of recent lattice conferences [129].

LQCD can be used, in principle, to simulate QCD at non-zero temperature and density, and in particular to study how confinement and chiral-symmetry breaking are lost as  $T$  and  $\mu$  (the chemical potential) are increased. This is of relevance to heavy-ion collisions, the early Universe and neutron-star structure. In practice, finite temperature simulations are computationally tractable and relatively mature, while simulations at finite  $\mu$  suffer from a “sign problem” and are at a rudimentary stage.

Another topic under active investigation is nucleon structure and inter-nucleon interactions. The simplest nucleon matrix elements are calculable with reasonable precision. Of particular interest are those of the axial current (leading to  $g_A$ ) and of the scalar density (with  $\langle N|\bar{s}s|N \rangle$  needed for dark matter searches).

Finally, we note that there is much recent interest in studying QCD-like theories with more fermions, possibly in other representations of the gauge group. The main interest is to find nearly conformal theories which might be candidates for “walking technicolor” models.

## 18.4. Outlook

While LQCD calculations have made major strides in the last decade, and are now playing an important role in constraining the Standard Model, there are many calculations that could be done in principle but are not yet mature due to limitations in computational resources. As we move to exascale resources (e.g.  $10^{18}$  floating point operations per second), the list of mature calculations will grow. Examples that we expect to mature in the next few years are results for excited hadrons, including quark-model exotics, at close to physical light-quark masses; results for moments of structure functions;  $K \rightarrow \pi\pi$  amplitudes (allowing a prediction of  $\epsilon'/\epsilon$  from the Standard Model);  $\bar{K} \leftrightarrow K$  and  $\bar{B} \leftrightarrow B$  mixing amplitudes from operators arising in models of new physics (allowing one to constrain these models in a manner complementary to the direct searches at the LHC); hadronic vacuum polarization contributions to muon  $g-2$ , the running of  $\alpha_{\text{EM}}$  and  $\alpha_s$ ;  $\pi \rightarrow \gamma\gamma$  and related amplitudes; long-distance contribution to  $\bar{K} \leftrightarrow K$  mixing and the light-by-light contribution to muon  $g-2$ . There will also be steady improvement in the precision attained for the mature quantities discussed above. As already noted, this will ultimately require simulations with  $m_u \neq m_d$  and including electromagnetic effects.

### References:

1. K.G. Wilson, Phys. Rev. **D10**, 2445 (1974).
2. T. Degrand & C. DeTar, “Lattice Methods for Quantum Chromodynamics,” World Scientific (2006).
3. C. Gattringer & C.B. Lang, “Quantum Chromodynamics on the Lattice: An Introductory Presentation,” Springer (2009).
4. “Modern Perspectives in Lattice QCD: quantum field theory and high performance computing” (Lecture notes of the Les Houches Summer School, Vol. 93) eds. L. Lellouch *et al.*, Oxford Univ. Press. (Aug. 2011).
5. W. Zimmermann, in “Lectures on Elementary Particles and Quantum Field Theory”, ed. S. Deser *et al.*, MIT Press, Cambridge, MA (1971); K. Symanzik, Nucl. Phys. **B226**, 187 (1983); Nucl. Phys. **B226**, 205 (1983).
6. M. Lüscher & P. Weisz, Commun. Math. Phys. **97**, 59 (1985).
7. Y. Iwasaki, UT-HEP-118.
8. H.B. Nielsen & M. Ninomiya, Phys. Lett. **B105**, 219 (1981).
9. B. Sheikholeslami & R. Wohlert, Nucl. Phys. **B259**, 572 (1985).
10. K. Jansen *et al.*, Phys. Lett. **B372**, 275 (1996).
11. M. Lüscher, JHEP **0305**, 052 (2003); Comp. Phys. Comm. **156**, 209 (2004) & Comp. Phys. Comm. **165**, 199 (2005);

- M. Hasenbusch, Phys. Lett. **B519**, 177 (2001);  
C. Urbach *et al.*, Comp. Phys. Comm. **174**, 87 (2006).
12. S. Durr *et al.*, Phys. Lett. **B705**, 477 (2011).
  13. N. Ishizuka *et al.*, Phys. Rev. **D92**, 074503 (2015).
  14. R. Frezzotti *et al.* [Alpha Collab.], JHEP **0108**, 058 (2001).
  15. R. Frezzotti & G.C. Rossi, JHEP **0408**, 007 (2004).
  16. L. Susskind, Phys. Rev. **D16**, 3031 (1977).
  17. M. Golterman, PoS **CONFINEMENT8**, 014 (2008).
  18. C. Bernard, Phys. Rev. **D73**, 114503 (2006);  
S.R. Sharpe, PoS **LAT 2006**, 022 (2006).
  19. G.P. Lepage, Phys. Rev. **D59**, 074502 (1999).
  20. A. Bazavov *et al.* [MILC Collab.], Rev. Mod. Phys. **82**, 1349 (2010).
  21. E. Follana *et al.* [HPQCD & UKQCD Collabs.], Phys. Rev. **D75**, 054502 (2007).
  22. P.H. Ginsparg & K.G. Wilson, Phys. Rev. **D25**, 2649 (1982).
  23. P. Hasenfratz *et al.*, Phys. Lett. **B427**, 125 (1998).
  24. M. Lüscher, Phys. Lett. **B428**, 342 (1998).
  25. D.B. Kaplan, Phys. Lett. **B288**, 342 (1992);  
Y. Shamir, Nucl. Phys. **B406**, 90 (1993);  
Y. Shamir, Nucl. Phys. **B417**, 167 (1994).
  26. H. Neuberger, Phys. Lett. **B417**, 141 (1998); Phys. Lett. **B427**, 353 (1998).
  27. A. Borici, hep-lat/9912040; A.D. Kennedy, hep-lat/060703.
  28. E. Shintani *et al.* [JLQCD Collab.], Phys. Rev. Lett. **101**, 242001 (2008).
  29. A. Bazavov *et al.* [MILC Collab.], Phys. Rev. **D87**, 054505 (2013).
  30. R. Baron *et al.* [ETM Collab.], JHEP **1006**, 111 (2010).
  31. E. Eichten & B. R. Hill, Phys. Lett. **B234**, 511 (1990).
  32. J. Heitger & R. Sommer [ALPHA Collab.], JHEP **0402**, 022 (2004);  
B. Blossier *et al.* [ALPHA Collab.], JHEP **1012**, 039 (2010).
  33. B.A. Thacker & G.P. Lepage, Phys. Rev. **D43**, 196 (1991);  
G.P. Lepage *et al.*, Phys. Rev. **D46**, 4052 (1992).
  34. A.X. El-Khadra *et al.*, Phys. Rev. **D55**, 3933 (1997).
  35. S. Aoki *et al.*, Prog. Theor. Phys. **109**, 383 (2003).
  36. N.H. Christ *et al.*, Phys. Rev. **D76**, 074505 (2007).
  37. Y. Aoki *et al.* [RBC and UKQCD Collabs.], Phys. Rev. **D86**, 116003 (2012).
  38. N.H. Christ *et al.*, Phys. Rev. **D91**, 054502 (2015).
  39. B. Blossier *et al.* [ETM Collab.], JHEP **1004**, 049 (2010).
  40. S. Borsanyi *et al.*, Science **347**, 1452 (2015).
  41. S. Aoki *et al.*, Eur. Phys. J. **C74**, 2890 (2014).
  42. S. Aoki *et al.* [PACS-CS Collab.], PTEP **2012**, 01A102 (2012).
  43. T. Ishikawa *et al.*, Phys. Rev. Lett. **109**, 072002 (2012).
  44. G.M. de Divitiis *et al.* [RM123 Collab.], Phys. Rev. **D87**, 114505 (2013).
  45. M. Bruno *et al.* [ALPHA Collab.], Phys. Rev. Lett. **114**, 102001 (2015).
  46. E. Follana *et al.* [HPQCD & UKQCD Collabs.], Phys. Rev. Lett. **100**, 062002 (2008);  
C.T.H. Davies *et al.* [HPQCD Collab.] Phys. Rev. **D82**, 114504 (2010).
  47. S. Schaefer *et al.* [ALPHA Collab.], Nucl. Phys. **B845**, 93 (2011).
  48. M. Lüscher, PoS **LATTICE 2010**, 015 (2010);  
S. Schaefer *et al.* [ALPHA Collab.], Nucl. Phys. **B845**, 93 (2011).
  49. R. Brower *et al.*, Phys. Lett. **B560**, 64 (2003).
  50. S. Aoki *et al.*, Phys. Rev. **D76**, 054508 (2007).
  51. G. Colangelo *et al.*, Nucl. Phys. **B721**, 136 (2005).
  52. M. Lüscher, Commun. Math. Phys. **104**, 177 (1986).
  53. G. Martinelli *et al.*, Nucl. Phys. **B445**, 81 (1995).
  54. M. Lüscher *et al.*, Nucl. Phys. **B384**, 168 (1992).
  55. M. Lüscher *et al.*, Nucl. Phys. **B413**, 481 (1994);  
M. Della Morte *et al.* [ALPHA Collab.], Nucl. Phys. **B713**, 378 (2005).
  56. S. Duane *et al.*, Phys. Lett. **B195**, 216 (1987).
  57. M.A. Clark & A.D. Kennedy, Phys. Rev. Lett. **98**, 051601 (2007).
  58. M. Creutz, Phys. Rev. **D38**, 1228 (1988);  
R. Gupta *et al.*, Phys. Rev. **D38**, 1278 (1988).
  59. J.E. Mandula *et al.*, Nucl. Phys. **B228**, 91 (1983);  
J.E. Mandula & E. Shpiz, Nucl. Phys. **B232**, 180 (1984).
  60. H.B. Meyer & M.J. Teper, Nucl. Phys. **B658**, 113 (2003).
  61. J.J. Dudek *et al.*, Phys. Rev. **D82**, 034508 (2010);  
J.J. Dudek *et al.*, Phys. Rev. **D83**, 111502 (2011);  
R.G. Edwards *et al.*, Phys. Rev. **D84**, 074508 (2011).
  62. M. Lüscher & U. Wolff, Nucl. Phys. **B339**, 222 (1990).
  63. G.P. Engel *et al.* [Bern-Graz-Regensburg Collab.], Phys. Rev. **D82**, 034505 (2010).
  64. M.S. Mahbub *et al.*, Ann. Phys. **342**, 270 (2014).
  65. M. Lüscher, Commun. Math. Phys. **105**, 153 (1986); Nucl. Phys. **B354**, 531 (1991), and Nucl. Phys. **B364**, 237 (1991).
  66. L. Maiani & M. Testa, Phys. Lett. **B245**, 585 (1990).
  67. S. Prelovsek *et al.*, Acta Phys. Polon. Supp. **6**, no. 3, 879 (2013);  
S. Prelovsek, PoS **LATTICE 2014**, 015 (2014).
  68. S.R. Beane *et al.*, Int. J. Mod. Phys. **E17**, 1157 (2008);  
M.J. Savage, Prog. in Part. Nucl. Phys. **67**, 140 (2012).
  69. K. Polejaeva & A. Rusetsky, Eur. Phys. J. **A48**, 67 (2012);  
R.A. Briceno & Z. Davoudi, Phys. Rev. **D87**, 094507 (2013);  
M.T. Hansen & S.R. Sharpe, Phys. Rev. **D90**, 116003 (2014);  
Phys. Rev. **D92**, 114509 (2015).
  70. L. Lellouch & M. Lüscher, Commun. Math. Phys. **219**, 31 (2001).
  71. T. Blum *et al.*, Phys. Rev. Lett. **108**, 141601 (2012); Phys. Rev. **D86**, 074513 (2012);  
Z. Bai *et al.*, Phys. Rev. Lett. **115**, 212001 (2015).
  72. T. Blum *et al.*, Phys. Rev. **D91**, 074502 (2015).
  73. B. Jager *et al.*, JHEP **1101**, 019 (2011);  
M. Doring *et al.*, Eur. Phys. J. **A47**, 139 (2011);  
M.T. Hansen & S.R. Sharpe, Phys. Rev. **D86**, 016007 (2012);  
R.A. Briceno & Z. Davoudi, Phys. Rev. **D88**, 094507 (2013);  
R.A. Briceno *et al.*, Phys. Rev. **D91**, 034501 (2015).
  74. X. Feng *et al.*, Phys. Rev. Lett. **107**, 081802 (2011).
  75. P. Boyle *et al.*, Phys. Rev. **D85**, 074504 (2012).
  76. M. Della Morte *et al.*, JHEP **1203**, 055 (2012).
  77. F. Burger *et al.* [ETM Collab.], JHEP **1402**, 099 (2014).
  78. T. Blum, Phys. Rev. Lett. **91**, 052001 (2003).
  79. T. Blum *et al.*, Phys. Rev. Lett. **114**, 012001 (2015).
  80. Z. Bai *et al.*, Phys. Rev. Lett. **113**, 112003 (2014).
  81. C. Bernard *et al.*, Nucl. Phys. (Proc. Supp.) **119**, 170 (2003).
  82. S. Perantonis & C. Michael, Nucl. Phys. **B347**, 854 (1990).
  83. G.S. Bali & K. Schilling, Phys. Rev. **D46**, 2636 (1992).
  84. S. Necco & R. Sommer, Nucl. Phys. **B622**, 328 (2002).
  85. J.M. Blairon *et al.*, Nucl. Phys. **B180**, 439 (1981).
  86. H. Fukaya *et al.* [JLQCD Collab.], Phys. Rev. Lett. **104**, 122002 (2010);  
H. Fukaya *et al.* [JLQCD & TWQCD Collabs.], Phys. Rev. **D83**, 074501 (2011).
  87. L. Giusti & M. Lüscher, JHEP **0903**, 013 (2009).
  88. K. Cichy *et al.*, JHEP **1310**, 175 (2013).
  89. G.P. Engel *et al.*, Phys. Rev. Lett. **114**, 112001 (2015).
  90. G.P. Engel *et al.*, Phys. Rev. **D91**, 054505 (2015).
  91. Z. Fodor & C. Hoelbling, Rev. Mod. Phys. **84**, 449 (2012).
  92. T. Blum *et al.* [RBC & UKQCD Collabs.], arXiv:1411.7017.
  93. T. Bae *et al.* [SWME Collab.], Phys. Rev. **D89**, 074504 (2014).
  94. J. Laiho & R.S. Van de Water, PoS **LATTICE 2011**, 293 (2011).
  95. E. Gamiz *et al.* [HPQCD Collab.], Phys. Rev. **D80**, 014503 (2009).
  96. C.M. Bouchard *et al.*, PoS **LATTICE 2011**, 274 (2011).
  97. Y. Aoki *et al.*, Phys. Rev. **D91**, 114505 (2015).
  98. M. Ademollo & R. Gatto, Phys. Rev. Lett. **13**, 264 (1964).
  99. H. Leutwyler & M. Roos, Z. Phys. **C25**, 91 (1984).
  100. P.A. Boyle *et al.*, Phys. Rev. Lett. **100**, 141601 (2008).
  101. V. Lubicz *et al.* [ETM Collab.], Phys. Rev. **D80**, 111502 (2009);  
PoS **LATTICE 2010**, 316 (2010).
  102. P.A. Boyle *et al.*, Eur. Phys. J. **C69**, 159 (2010).

103. A. Bazavov *et al.* [FNAL/MILC Collabs.], Phys. Rev. **D87**, 073012 (2013).
104. T. Kaneko *et al.* [JLQCD Collab.], PoS LATTICE **2012**, 111 (2012).
105. P.A. Boyle *et al.*, JHEP **1308**, 132 (2013).
106. P.A. Boyle *et al.* [RBC/UKQCD Collab.], JHEP **1506**, 164 (2015).
107. H. Na *et al.* [HPQCD Collab.], Phys. Rev. **D84**, 114505 (2011).
108. H. Na *et al.* [HPQCD Collab.], Phys. Rev. **D82**, 114506 (2010).
109. E. Dalgic *et al.* [HPQCD Collab.], Phys. Rev. **D73**, 074502 (2006).
110. J.M. Flynn *et al.* [RBC/UKQCD Collabs.], Phys. Rev. **D91**, 074510 (2015).
111. J.A. Bailey *et al.* [Fermilab Lattice & MILC Collabs.], Phys. Rev. **D92**, 014024 (2015).
112. C. Bourrely *et al.*, Nucl. Phys. **B189**, 157 (1981);  
C.G. Boyd *et al.*, Phys. Rev. Lett. **74**, 4603 (1995);  
T. Becher & R.J. Hill, Phys. Lett. **B633**, 61 (2006);  
C. Bourrely *et al.*, Phys. Rev. **D79**, 013008 (2009).
113. M.C. Arnesen *et al.*, Phys. Rev. Lett. **95**, 071802 (2005).
114. J.A. Bailey *et al.*, Phys. Rev. **D79**, 054507 (2009).
115. C. Bernard *et al.*, Phys. Rev. **D79**, 014506 (2009);;  
J.A. Bailey *et al.* [Fermilab Lattice & MILC Collabs.], PoS **LATTICE2010**, 311 (2010).
116. J.A. Bailey *et al.* [Fermilab Lattice & MILC Collabs.], Phys. Rev. **D89**, 114504 (2014).
117. J.A. Bailey *et al.* [Fermilab Lattice & MILC Collabs.], to appear in Phys. Rev. **D**, (2016).
118. C.T. H. Davies *et al.* [HPQCD Collab.], Phys. Rev. **D78**, 114507 (2008).
119. E. Shintani *et al.*, Phys. Rev. **D82**, 074505 (2010).
120. I. Allison *et al.* [HPQCD Collab.], Phys. Rev. **D78**, 054513 (2008).
121. C. McNeile *et al.* [HPQCD Collab.], Phys. Rev. **D82**, 034512 (2010).
122. B. Chakraborty *et al.*, Phys. Rev. **D91**, 054508 (2015).
123. A. Bazavov *et al.*, Phys. Rev. **D86**, 114031 (2012).
124. B. Blossier *et al.*, Phys. Rev. **D85**, 034503 (2012); Phys. Rev. Lett. **108**, 262002 (2012).
125. S. Aoki *et al.* [PACS-CS Collab.], JHEP **0910**, 053 (2009).
126. P. Fritzsch *et al.*, PoS LATTICE **2014**, 291 (2014).
127. S. Capitani *et al.* [ALPHA Collab.], Nucl. Phys. **B544**, 669 (1999).
128. A. Bussone *et al.*, [arXiv:1411.0484](https://arxiv.org/abs/1411.0484).
129. N. Christ *et al.* (ed.), PoS Lattice **2014** (2014).

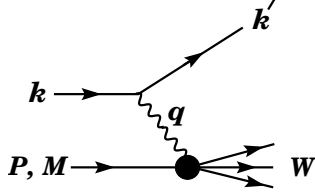


## 19. STRUCTURE FUNCTIONS

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### 19.1. Deep inelastic scattering

High-energy lepton-nucleon scattering (deep inelastic scattering) plays a key role in determining the partonic structure of the proton. The process  $\ell N \rightarrow \ell' X$  is illustrated in Fig. 19.1. The filled circle in this figure represents the internal structure of the proton which can be expressed in terms of structure functions.



**Figure 19.1:** Kinematic quantities for the description of deep inelastic scattering. The quantities  $k$  and  $k'$  are the four-momenta of the incoming and outgoing leptons,  $P$  is the four-momentum of a nucleon with mass  $M$ , and  $W$  is the mass of the recoiling system  $X$ . The exchanged particle is a  $\gamma$ ,  $W^\pm$ , or  $Z$ ; it transfers four-momentum  $q = k - k'$  to the nucleon.

Invariant quantities:

$\nu = \frac{q \cdot P}{M} = E - E'$  is the lepton's energy loss in the nucleon rest frame (in earlier literature sometimes  $\nu = q \cdot P$ ). Here,  $E$  and  $E'$  are the initial and final lepton energies in the nucleon rest frame.

$Q^2 = -q^2 = 2(E E' - \vec{k} \cdot \vec{k}') - m_\ell^2 - m_{\ell'}^2$  where  $m_\ell(m_{\ell'})$  is the initial (final) lepton mass. If  $E E' \sin^2(\theta/2) \gg m_\ell^2, m_{\ell'}^2$ , then

$\approx 4 E E' \sin^2(\theta/2)$ , where  $\theta$  is the lepton's scattering angle with respect to the lepton beam direction.

$x = \frac{Q^2}{2M\nu}$  where, in the parton model,  $x$  is the fraction of the nucleon's momentum carried by the struck quark.

$y = \frac{q \cdot P}{k \cdot P} = \frac{\nu}{E}$  is the fraction of the lepton's energy lost in the nucleon rest frame.

$W^2 = (P + q)^2 = M^2 + 2M\nu - Q^2$  is the mass squared of the system  $X$  recoiling against the scattered lepton.

$s = (k + P)^2 = \frac{Q^2}{xy} + M^2 + m_\ell^2$  is the center-of-mass energy squared of the lepton-nucleon system.

The process in Fig. 19.1 is called deep ( $Q^2 \gg M^2$ ) inelastic ( $W^2 \gg M^2$ ) scattering (DIS). In what follows, the masses of the initial and scattered leptons,  $m_\ell$  and  $m_{\ell'}$ , are neglected.

#### 19.1.1. DIS cross sections :

The double-differential cross section for deep inelastic scattering can be expressed in terms of kinematic variables in several ways.

$$\frac{d^2\sigma}{dx dy} = x(s - M^2) \frac{d^2\sigma}{dx dQ^2} = \frac{2\pi M\nu}{E'} \frac{d^2\sigma}{d\Omega_{N\text{rest}} dE'} . \quad (19.1)$$

In lowest-order perturbation theory, the cross section for the scattering of polarized leptons on polarized nucleons can be expressed in terms of the products of leptonic and hadronic tensors associated with the coupling of the exchanged bosons at the upper and lower vertices in Fig. 19.1 (see Refs. 1–4)

$$\frac{d^2\sigma}{dx dy} = \frac{2\pi y \alpha^2}{Q^4} \sum_j \eta_j L_j^{\mu\nu} W_{\mu\nu}^j . \quad (19.2)$$

For neutral-current processes, the summation is over  $j = \gamma, Z$  and  $\gamma Z$  representing photon and  $Z$  exchange and the interference between them, whereas for charged-current interactions there is only  $W$  exchange,  $j = W$ . (For transverse nucleon polarization, there is a dependence on the azimuthal angle of the scattered lepton.) The lepton tensor  $L_{\mu\nu}$  is associated with the coupling of the exchange boson to the leptons. For incoming leptons of charge  $e = \pm 1$  and helicity  $\lambda = \pm 1$ ,

$$\begin{aligned} L_{\mu\nu}^\gamma &= 2 \left( k_\mu k'_\nu + k'_\mu k_\nu - (k \cdot k' - m_\ell^2) g_{\mu\nu} - i \lambda \varepsilon_{\mu\nu\alpha\beta} k^\alpha k'^\beta \right), \\ L_{\mu\nu}^{\gamma Z} &= (g_V^e + e \lambda g_A^e) L_{\mu\nu}^\gamma, \quad L_{\mu\nu}^Z = (g_V^e + e \lambda g_A^e)^2 L_{\mu\nu}^\gamma, \\ L_{\mu\nu}^W &= (1 + e \lambda)^2 L_{\mu\nu}^\gamma, \end{aligned} \quad (19.3)$$

where  $g_V^e = -\frac{1}{2} + 2 \sin^2 \theta_W$ ,  $g_A^e = -\frac{1}{2}$ .

Although here the helicity formalism is adopted, an alternative approach is to express the tensors in Eq. (19.3) in terms of the polarization of the lepton.

The factors  $\eta_j$  in Eq. (19.2) denote the ratios of the corresponding propagators and couplings to the photon propagator and coupling squared

$$\begin{aligned} \eta_\gamma &= 1 \quad ; \quad \eta_{\gamma Z} = \left( \frac{G_F M_Z^2}{2\sqrt{2}\pi\alpha} \right) \left( \frac{Q^2}{Q^2 + M_Z^2} \right); \\ \eta_Z &= \eta_{\gamma Z}^2 \quad ; \quad \eta_W = \frac{1}{2} \left( \frac{G_F M_W^2}{4\pi\alpha} \frac{Q^2}{Q^2 + M_W^2} \right)^2. \end{aligned} \quad (19.4)$$

The hadronic tensor, which describes the interaction of the appropriate electroweak currents with the target nucleon, is given by

$$W_{\mu\nu} = \frac{1}{4\pi} \int d^4z e^{iq \cdot z} \langle P, S | [J_\mu^\dagger(z), J_\nu(0)] | P, S \rangle, \quad (19.5)$$

where  $S$  denotes the nucleon-spin 4-vector, with  $S^2 = -M^2$  and  $S \cdot P = 0$ .

### 19.2. Structure functions of the proton

The structure functions are defined in terms of the hadronic tensor (see Refs. 1–3)

$$\begin{aligned} W_{\mu\nu} &= \left( -g_{\mu\nu} + \frac{q_\mu q_\nu}{q^2} \right) F_1(x, Q^2) + \frac{\hat{P}_\mu \hat{P}_\nu}{P \cdot q} F_2(x, Q^2) \\ &\quad - i \varepsilon_{\mu\nu\alpha\beta} \frac{q^\alpha P^\beta}{2P \cdot q} F_3(x, Q^2) \\ &\quad + i \varepsilon_{\mu\nu\alpha\beta} \frac{q^\alpha}{P \cdot q} \left[ S^\beta g_1(x, Q^2) + \left( S^\beta - \frac{S \cdot q}{P \cdot q} P^\beta \right) g_2(x, Q^2) \right] \\ &\quad + \frac{1}{P \cdot q} \left[ \frac{1}{2} (\hat{P}_\mu \hat{S}_\nu + \hat{S}_\mu \hat{P}_\nu) - \frac{S \cdot q}{P \cdot q} \hat{P}_\mu \hat{P}_\nu \right] g_3(x, Q^2) \\ &\quad + \frac{S \cdot q}{P \cdot q} \left[ \frac{\hat{P}_\mu \hat{P}_\nu}{P \cdot q} g_4(x, Q^2) + \left( -g_{\mu\nu} + \frac{q_\mu q_\nu}{q^2} \right) g_5(x, Q^2) \right] \end{aligned} \quad (19.6)$$

where

$$\hat{P}_\mu = P_\mu - \frac{P \cdot q}{q^2} q_\mu, \quad \hat{S}_\mu = S_\mu - \frac{S \cdot q}{q^2} q_\mu. \quad (19.7)$$

In Ref. 2, the definition of  $W_{\mu\nu}$  with  $\mu \leftrightarrow \nu$  is adopted, which changes the sign of the  $\varepsilon_{\mu\nu\alpha\beta}$  terms in Eq. (19.6), although the formulae given below are unchanged. Ref. 1 tabulates the relation between the structure functions defined in Eq. (19.6) and other choices available in the literature.

The cross sections for neutral- and charged-current deep inelastic scattering on unpolarized nucleons can be written in terms of the structure functions in the generic form

$$\begin{aligned} \frac{d^2\sigma^i}{dx dy} &= \frac{4\pi\alpha^2}{xyQ^2} \eta^i \left\{ \left( 1 - y - \frac{x^2 y^2 M^2}{Q^2} \right) F_2^i \right. \\ &\quad \left. + y^2 x F_1^i \mp \left( y - \frac{y^2}{2} \right) x F_3^i \right\}, \end{aligned} \quad (19.8)$$

where  $i = \text{NC, CC}$  corresponds to neutral-current ( $eN \rightarrow eX$ ) or charged-current ( $eN \rightarrow \nu X$  or  $\nu N \rightarrow eX$ ) processes, respectively. For incoming neutrinos,  $L_{\mu\nu}^W$  of Eq. (19.3) is still true, but with  $e, \lambda$  corresponding to the outgoing charged lepton. In the last term of Eq. (19.8), the  $-$  sign is taken for an incoming  $e^+$  or  $\bar{\nu}$  and the  $+$  sign for an incoming  $e^-$  or  $\nu$ . The factor  $\eta^{\text{NC}} = 1$  for unpolarized  $e^\pm$  beams, whereas\*

$$\eta^{\text{CC}} = (1 \pm \lambda)^2 \eta_W \quad (19.9)$$

with  $\pm$  for  $\ell^\pm$ ; and where  $\lambda$  is the helicity of the incoming lepton and  $\eta_W$  is defined in Eq. (19.4); for incoming neutrinos  $\eta^{\text{CC}} = 4\eta_W$ . The CC structure functions, which derive exclusively from  $W$  exchange, are

$$F_1^{\text{CC}} = F_1^W, \quad F_2^{\text{CC}} = F_2^W, \quad xF_3^{\text{CC}} = xF_3^W. \quad (19.10)$$

The NC structure functions  $F_2^\gamma, F_2^{\gamma Z}, F_2^Z$  are, for  $e^\pm N \rightarrow e^\pm X$ , given by Ref. 5,

$$F_2^{\text{NC}} = F_2^\gamma - (g_V^e \pm \lambda g_A^e) \eta_{\gamma Z} F_2^{\gamma Z} + (g_V^e \pm g_A^e \pm 2\lambda g_V^e g_A^e) \eta_Z F_2^Z \quad (19.11)$$

and similarly for  $F_1^{\text{NC}}$ , whereas

$$xF_3^{\text{NC}} = -(g_A^e \pm \lambda g_V^e) \eta_{\gamma Z} xF_3^{\gamma Z} + [2g_V^e g_A^e \pm \lambda(g_V^e \pm g_A^e)] \eta_Z xF_3^Z. \quad (19.12)$$

The polarized cross-section difference

$$\Delta\sigma = \sigma(\lambda_n = -1, \lambda_\ell) - \sigma(\lambda_n = 1, \lambda_\ell), \quad (19.13)$$

where  $\lambda_\ell, \lambda_n$  are the helicities ( $\pm 1$ ) of the incoming lepton and nucleon, respectively, may be expressed in terms of the five structure functions  $g_{1,\dots,5}(x, Q^2)$  of Eq. (19.6). Thus,

$$\begin{aligned} \frac{d^2 \Delta\sigma^i}{dx dy} &= \frac{8\pi\alpha^2}{xyQ^2} \eta^i \left\{ -\lambda_\ell y \left( 2 - y - 2x^2 y^2 \frac{M^2}{Q^2} \right) x g_1^i + \lambda_\ell 4x^3 y^2 \frac{M^2}{Q^2} g_2^i \right. \\ &+ 2x^2 y \frac{M^2}{Q^2} \left( 1 - y - x^2 y^2 \frac{M^2}{Q^2} \right) g_3^i \\ &\left. - \left( 1 + 2x^2 y \frac{M^2}{Q^2} \right) \left[ \left( 1 - y - x^2 y^2 \frac{M^2}{Q^2} \right) g_4^i + xy^2 g_5^i \right] \right\} \quad (19.14) \end{aligned}$$

with  $i = \text{NC or CC}$  as before. The Eq. (19.13) corresponds to the difference of antiparallel minus parallel spins of the incoming particles for  $e^-$  or  $\nu$  initiated reactions, but the difference of parallel minus antiparallel for  $e^+$  or  $\bar{\nu}$  initiated processes. For longitudinal nucleon polarization, the contributions of  $g_2$  and  $g_3$  are suppressed by powers of  $M^2/Q^2$ . These structure functions give an unsuppressed contribution to the cross section for transverse polarization [1], but in this case the cross-section difference vanishes as  $M/Q \rightarrow 0$ .

Because the same tensor structure occurs in the spin-dependent and spin-independent parts of the hadronic tensor of Eq. (19.6) in the  $M^2/Q^2 \rightarrow 0$  limit, the differential cross-section difference of Eq. (19.14) may be obtained from the differential cross section of Eq. (19.8) by replacing

$$F_1 \rightarrow -g_5, \quad F_2 \rightarrow -g_4, \quad F_3 \rightarrow 2g_1, \quad (19.15)$$

and multiplying by two, since the total cross section is the average over the initial-state polarizations. In this limit, Eq. (19.8) and Eq. (19.14) may be written in the form

$$\begin{aligned} \frac{d^2 \sigma^i}{dx dy} &= \frac{2\pi\alpha^2}{xyQ^2} \eta^i \left[ Y_+ F_2^i \mp Y_- x F_3^i - y^2 F_L^i \right], \\ \frac{d^2 \Delta\sigma^i}{dx dy} &= \frac{4\pi\alpha^2}{xyQ^2} \eta^i \left[ -Y_+ g_4^i \mp Y_- 2x g_1^i + y^2 g_L^i \right], \quad (19.16) \end{aligned}$$

with  $i = \text{NC or CC}$ , where  $Y_\pm = 1 \pm (1 - y)^2$  and

$$F_L^i = F_2^i - 2xF_1^i, \quad g_L^i = g_4^i - 2xg_5^i. \quad (19.17)$$

In the naive quark-parton model, the analogy with the Callan-Gross relations [6]  $F_L^i = 0$ , are the Dicus relations [7]  $g_L^i = 0$ . Therefore, there are only two independent polarized structure functions:  $g_1$  (parity conserving) and  $g_5$  (parity violating), in analogy with the unpolarized structure functions  $F_1$  and  $F_3$ .

### 19.2.1. Structure functions in the quark-parton model :

In the quark-parton model [8,9], contributions to the structure functions  $F^i$  and  $g^i$  can be expressed in terms of the quark distribution functions  $q(x, Q^2)$  of the proton, where  $q = u, \bar{u}, d, \bar{d}$  etc. The quantity  $q(x, Q^2)dx$  is the number of quarks (or antiquarks) of designated flavor that carry a momentum fraction between  $x$  and  $x + dx$  of the proton's momentum in a frame in which the proton momentum is large.

For the neutral-current processes  $ep \rightarrow eX$ ,

$$\begin{aligned} [F_2^\gamma, F_2^{\gamma Z}, F_2^Z] &= x \sum_q [e_q^2, 2e_q g_V^q, g_V^{q^2} + g_A^{q^2}] (q + \bar{q}), \\ [F_3^\gamma, F_3^{\gamma Z}, F_3^Z] &= \sum_q [0, 2e_q g_A^q, 2g_V^q g_A^q] (q - \bar{q}), \\ [g_1^\gamma, g_1^{\gamma Z}, g_1^Z] &= \frac{1}{2} \sum_q [e_q^2, 2e_q g_V^q, g_V^{q^2} + g_A^{q^2}] (\Delta q + \Delta \bar{q}), \\ [g_5^\gamma, g_5^{\gamma Z}, g_5^Z] &= \sum_q [0, e_q g_A^q, g_V^q g_A^q] (\Delta q - \Delta \bar{q}), \quad (19.18) \end{aligned}$$

where  $g_V^q = \pm \frac{1}{2} - 2e_q \sin^2 \theta_W$  and  $g_A^q = \pm \frac{1}{2}$ , with  $\pm$  according to whether  $q$  is a  $u$ - or  $d$ -type quark respectively. The quantity  $\Delta q$  is the difference  $q \uparrow - q \downarrow$  of the distributions with the quark spin parallel and antiparallel to the proton spin.

For the charged-current processes  $e^- p \rightarrow \nu X$  and  $\bar{\nu} p \rightarrow e^+ X$ , the structure functions are:

$$\begin{aligned} F_2^{W-} &= 2x(u + \bar{d} + \bar{s} + c \dots), \\ F_3^{W-} &= 2(u - \bar{d} - \bar{s} + c \dots), \\ g_1^{W-} &= (\Delta u + \Delta \bar{d} + \Delta \bar{s} + \Delta c \dots), \\ g_5^{W-} &= (-\Delta u + \Delta \bar{d} + \Delta \bar{s} - \Delta c \dots), \quad (19.19) \end{aligned}$$

where only the active flavors have been kept and where CKM mixing has been neglected. For  $e^+ p \rightarrow \bar{\nu} X$  and  $\nu p \rightarrow e^- X$ , the structure functions  $F^{W+}, g^{W+}$  are obtained by the flavor interchanges  $d \leftrightarrow u, s \leftrightarrow c$  in the expressions for  $F^{W-}, g^{W-}$ . The structure functions for scattering on a neutron are obtained from those of the proton by the interchange  $u \leftrightarrow d$ . For both the neutral- and charged-current processes, the quark-parton model predicts  $2xF_1^i = F_2^i$  and  $g_4^i = 2xg_5^i$ .

Neglecting masses, the structure functions  $g_2$  and  $g_3$  contribute only to scattering from transversely polarized nucleons (for which  $S \cdot q = 0$ ), and have no simple interpretation in terms of the quark-parton model. They arise from off-diagonal matrix elements  $\langle P, \lambda' | [J_\mu^\dagger(z), J_\nu(0)] | P, \lambda \rangle$ , where the proton helicities satisfy  $\lambda' \neq \lambda$ . In fact, the leading-twist contributions to both  $g_2$  and  $g_3$  are both twist-2 and twist-3, which contribute at the same order of  $Q^2$ . The Wandzura-Wilczek relation [10] expresses the twist-2 part of  $g_2$  in terms of  $g_1$  as

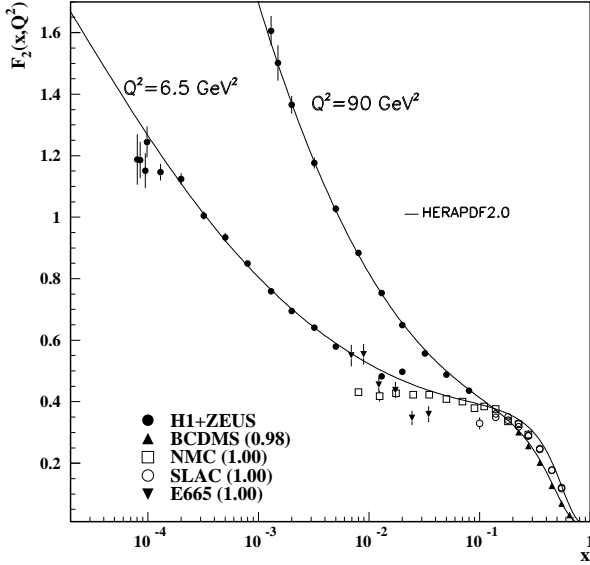
$$g_2^i(x) = -g_1^i(x) + \int_x^1 \frac{dy}{y} g_1^i(y). \quad (19.20)$$

However, the twist-3 component of  $g_2$  is unknown. Similarly, there is a relation expressing the twist-2 part of  $g_3$  in terms of  $g_4$ . A complete set of relations, including  $M^2/Q^2$  effects, can be found in Ref. 11.

### 19.2.2. Structure functions and QCD :

One of the most striking predictions of the quark-parton model is that the structure functions  $F_i, g_i$  scale, i.e.,  $F_i(x, Q^2) \rightarrow F_i(x)$  in the Bjorken limit that  $Q^2$  and  $\nu \rightarrow \infty$  with  $x$  fixed [12]. This property is related to the assumption that the transverse momentum of the partons in the infinite-momentum frame of the proton is small. In QCD, however, the radiation of hard gluons from the quarks violates this assumption, leading to logarithmic scaling violations, which are particularly large at small  $x$ , see Fig. 19.2. The radiation of gluons produces the evolution of the structure functions. As  $Q^2$  increases, more and more gluons are radiated, which in turn split into  $q\bar{q}$  pairs. This process leads both to the softening of the initial quark momentum

distributions and to the growth of the gluon density and the  $q\bar{q}$  sea as  $x$  decreases. For spin-dependent structure functions, data exists for a more restricted range of  $Q^2$  and has lower precision, so that the scaling violations are not seen so clearly. However, spin-dependent parton distributions have been extracted by comparison to data; Fig. 19.3 shows several versions (discussed in more detail in Sec. 19.3 below) at a scale of  $2.5 \text{ GeV}^2$  compared to the data from semi-inclusive DIS.



**Figure 19.2:** The proton structure function  $F_2^p$  given at two  $Q^2$  values ( $6.5 \text{ GeV}^2$  and  $90 \text{ GeV}^2$ ), which exhibit scaling at the ‘pivot’ point  $x \sim 0.14$ . See the captions in Fig. 19.8 and Fig. 19.10 for the references of the data. The various data sets have been renormalized by the factors shown in brackets in the key to the plot, which were globally determined in a previous HERAPDF analysis [13]. The curves were obtained using the PDFs from the HERAPDF analysis [14]. In practice, data for the reduced cross section,  $F_2(x, Q^2) - (y^2/Y_+)F_L(x, Q^2)$ , are fitted, rather than  $F_2$  and  $F_L$  separately.

In QCD, the above processes are described in terms of scale-dependent parton distributions  $f_a(x, \mu^2)$ , where  $a = g$  or  $q$  and, typically,  $\mu$  is the scale of the probe  $Q$ . For  $Q^2 \gg M^2$ , the structure functions are of the form

$$F_i = \sum_a C_i^a \otimes f_a, \quad (19.21)$$

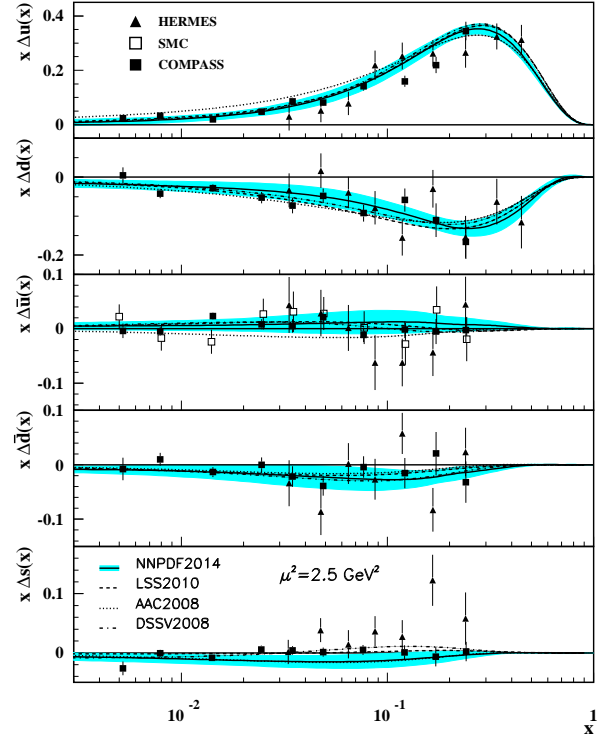
where  $\otimes$  denotes the convolution integral

$$C \otimes f = \int_x^1 \frac{dy}{y} C(y) f\left(\frac{x}{y}\right), \quad (19.22)$$

and where the coefficient functions  $C_i^a$  are given as a power series in  $\alpha_s$ . The parton distribution  $f_a$  corresponds, at a given  $x$ , to the density of parton  $a$  in the proton integrated over transverse momentum  $k_t$  up to  $\mu$ . Its evolution in  $\mu$  is described in QCD by a DGLAP equation (see Refs. 24–27) which has the schematic form

$$\frac{\partial f_a}{\partial \ln \mu^2} \sim \frac{\alpha_s(\mu^2)}{2\pi} \sum_b (P_{ab} \otimes f_b), \quad (19.23)$$

where the  $P_{ab}$ , which describe the parton splitting  $b \rightarrow a$ , are also given as a power series in  $\alpha_s$ . Although perturbative QCD can predict, via Eq. (19.23), the evolution of the parton distribution functions from a particular scale,  $\mu_0$ , these DGLAP equations cannot predict them *a priori* at any particular  $\mu_0$ . Thus they must be measured at a starting point  $\mu_0$  before the predictions of QCD can be compared to the data at other scales,  $\mu$ . In general, all observables involving a hard



**Figure 19.3:** Distributions of  $x$  times the polarized parton distributions  $\Delta q(x)$  (where  $q = u, d, \bar{u}, \bar{d}, s$ ) using the NNPDF2014 [15], AAC2008 [16], DSSV2008 [17], and LSS2010 [18] parameterizations at a scale  $\mu^2 = 2.5 \text{ GeV}^2$ , showing the blue-shaded error corridor of the NNPDF2014 set. The points represent data from semi-inclusive positron (HERMES [19,20]) and muon (SMC [21] and COMPASS [22,23]) deep inelastic scattering given at  $Q^2 = 2.5 \text{ GeV}^2$ . The SMC results are extracted under the assumption that  $\Delta \bar{u}(x) = \Delta \bar{d}(x)$ .

hadronic interaction (such as structure functions) can be expressed as a convolution of calculable, process-dependent coefficient functions and these universal parton distributions, e.g. Eq. (19.21).

It is often convenient to write the evolution equations in terms of the gluon, non-singlet ( $q^{NS}$ ) and singlet ( $q^S$ ) quark distributions, such that

$$q^{NS} = q_i - \bar{q}_i \quad (\text{or } q_i - q_j), \quad q^S = \sum_i (q_i + \bar{q}_i). \quad (19.24)$$

The non-singlet distributions have non-zero values of flavor quantum numbers, such as isospin and baryon number. The DGLAP evolution equations then take the form

$$\begin{aligned} \frac{\partial q^{NS}}{\partial \ln \mu^2} &= \frac{\alpha_s(\mu^2)}{2\pi} P_{qq} \otimes q^{NS}, \\ \frac{\partial}{\partial \ln \mu^2} \begin{pmatrix} q^S \\ g \end{pmatrix} &= \frac{\alpha_s(\mu^2)}{2\pi} \begin{pmatrix} P_{qq} & 2n_f P_{qg} \\ P_{gq} & P_{gg} \end{pmatrix} \otimes \begin{pmatrix} q^S \\ g \end{pmatrix}, \end{aligned} \quad (19.25)$$

where  $P$  are splitting functions that describe the probability of a given parton splitting into two others, and  $n_f$  is the number of (active) quark flavors. The leading-order Altarelli-Parisi [26] splitting functions are

$$P_{qq} = \frac{4}{3} \left[ \frac{1+x^2}{(1-x)} \right]_+ = \frac{4}{3} \left[ \frac{1+x^2}{(1-x)_+} \right] + 2\delta(1-x), \quad (19.26)$$

$$P_{gg} = \frac{1}{2} \left[ x^2 + (1-x)^2 \right], \quad (19.27)$$

$$P_{gq} = \frac{4}{3} \left[ \frac{1+(1-x)^2}{x} \right], \quad (19.28)$$

$$P_{gg} = 6 \left[ \frac{1-x}{x} + x(1-x) + \frac{x}{(1-x)_+} \right] + \left[ \frac{11}{2} - \frac{n_f}{3} \right] \delta(1-x), \quad (19.29)$$

where the notation  $[F(x)]_+$  defines a distribution such that for any sufficiently regular test function,  $f(x)$ ,

$$\int_0^1 dx f(x) [F(x)]_+ = \int_0^1 dx (f(x) - f(1)) F(x). \quad (19.30)$$

In general, the splitting functions can be expressed as a power series in  $\alpha_s$ . The series contains both terms proportional to  $\ln \mu^2$  and to  $\ln(1/x)$  and  $\ln(1-x)$ . The leading-order DGLAP evolution sums up the  $(\alpha_s \ln \mu^2)^n$  contributions, while at next-to-leading order (NLO) the sum over the  $\alpha_s (\alpha_s \ln \mu^2)^{n-1}$  terms is included [28,29]. The NNLO contributions to the splitting functions and the DIS coefficient functions are also all known [30–32].

In the kinematic region of very small  $x$ , one may also sum leading terms in  $\ln(1/x)$ , independent of the value of  $\ln \mu^2$ . At leading order, LL $x$ , this is done by the BFKL equation for the unintegrated distributions (see Refs. [33,34]). The leading-order  $(\alpha_s \ln(1/x))^n$  terms result in a power-like growth,  $x^{-\omega}$  with  $\omega = (12\alpha_s \ln 2)/\pi$ , at asymptotic values of  $\ln 1/x$ . The next-to-leading  $\ln 1/x$  (NLL $x$ ) contributions are also available [35,36]. They are so large (and negative) that the results initially appeared to be perturbatively unstable. Methods, based on a combination of collinear and small- $x$  resummations, have been developed which reorganize the perturbative series into a more stable hierarchy [37–40]. There are some limited indications that small- $x$  resummations become necessary for sufficient precision for  $x \lesssim 10^{-3}$  at low scales. There is not yet any very convincing indication for a ‘non-linear’ regime, for  $Q^2 \gtrsim 2 \text{ GeV}^2$ , in which the gluon density would be so high that gluon-gluon recombination effects would become significant.

**Table 19.1:** The main processes relevant to global PDF analyses, ordered in three groups: fixed-target experiments, HERA and the  $p\bar{p}$  Tevatron /  $pp$  LHC. For each process we give an indication of their dominant partonic subprocesses, the primary partons which are probed and the approximate range of  $x$  constrained by the data.

Process	Subprocess	Partons	$x$ range
$\ell^\pm \{p, n\} \rightarrow \ell^\pm X$	$\gamma^* q \rightarrow q$	$q, \bar{q}, g$	$x \gtrsim 0.01$
$\ell^\pm n/p \rightarrow \ell^\pm X$	$\gamma^* d/u \rightarrow d/u$	$d/u$	$x \gtrsim 0.01$
$pp \rightarrow \mu^+ \mu^- X$	$u\bar{u}, d\bar{d} \rightarrow \gamma^*$	$\bar{q}$	$0.015 \lesssim x \lesssim 0.35$
$pn/pp \rightarrow \mu^+ \mu^- X$	$(u\bar{d})/(u\bar{u}) \rightarrow \gamma^*$	$\bar{d}/\bar{u}$	$0.015 \lesssim x \lesssim 0.35$
$\nu(\bar{\nu}) N \rightarrow \mu^-(\mu^+) X$	$W^* q \rightarrow q'$	$q, \bar{q}$	$0.01 \lesssim x \lesssim 0.5$
$\nu N \rightarrow \mu^- \mu^+ X$	$W^* s \rightarrow c$	$s$	$0.01 \lesssim x \lesssim 0.2$
$\bar{\nu} N \rightarrow \mu^+ \mu^- X$	$W^* \bar{s} \rightarrow \bar{c}$	$\bar{s}$	$0.01 \lesssim x \lesssim 0.2$
$e^\pm p \rightarrow e^\pm X$	$\gamma^* q \rightarrow q$	$g, q, \bar{q}$	$10^{-4} \lesssim x \lesssim 0.1$
$e^+ p \rightarrow \bar{\nu} X$	$W^+ \{d, s\} \rightarrow \{u, c\}$	$d, s$	$x \gtrsim 0.01$
$e^\pm p \rightarrow e^\pm c\bar{c}X, e^\pm b\bar{b}X$	$\gamma^* c \rightarrow c, \gamma^* g \rightarrow c\bar{c}$	$c, b, g$	$10^{-4} \lesssim x \lesssim 0.01$
$e^\pm p \rightarrow \text{jet} + X$	$\gamma^* g \rightarrow q\bar{q}$	$g$	$0.01 \lesssim x \lesssim 0.1$
$p\bar{p}, pp \rightarrow \text{jet} + X$	$gg, qg, q\bar{q} \rightarrow 2j$	$g, q$	$0.00005 \lesssim x \lesssim 0.5$
$p\bar{p} \rightarrow (W^\pm \rightarrow \ell^\pm \nu) X$	$ud \rightarrow W^+, \bar{u}\bar{d} \rightarrow W^-$	$u, d, \bar{u}, \bar{d}$	$x \gtrsim 0.05$
$pp \rightarrow (W^\pm \rightarrow \ell^\pm \nu) X$	$u\bar{d} \rightarrow W^+, d\bar{u} \rightarrow W^-$	$u, d, \bar{u}, \bar{d}, g$	$x \gtrsim 0.001$
$p\bar{p}(pp) \rightarrow (Z \rightarrow \ell^+ \ell^-) X$	$uu, dd, ..(u\bar{u}, ..) \rightarrow Z$	$u, d, ..(g)$	$x \gtrsim 0.001$
$pp \rightarrow W^- c, W^+ \bar{c}$	$gs \rightarrow W^- c$	$s, \bar{s}$	$x \sim 0.01$
$pp \rightarrow (\gamma^* \rightarrow \ell^+ \ell^-) X$	$u\bar{u}, d\bar{d}, .. \rightarrow \gamma^*$	$\bar{q}, g$	$x \gtrsim 10^{-5}$
$pp \rightarrow b\bar{b} X, t\bar{t} X$	$gg \rightarrow b\bar{b}, t\bar{t}$	$g$	$x \gtrsim 10^{-5}, 10^{-2}$
$pp \rightarrow \text{exclusive } J/\psi, \Upsilon$	$\gamma^*(gg) \rightarrow J/\psi, \Upsilon$	$g$	$x \gtrsim 10^{-5}, 10^{-4}$
$pp \rightarrow \gamma X$	$gq \rightarrow \gamma q, g\bar{q} \rightarrow \gamma \bar{q}$	$g$	$x \gtrsim 0.005$

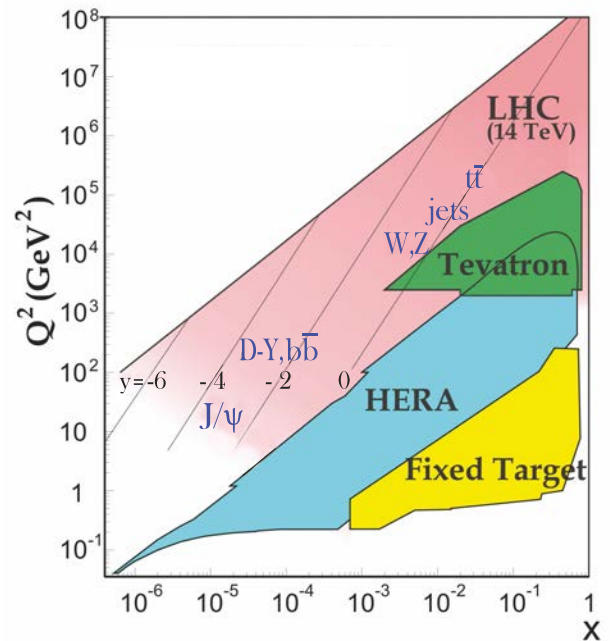
The precision of the experimental data demands that at least NLO, and preferably NNLO, DGLAP evolution be used in comparisons between QCD theory and experiment. Beyond the leading order, it is necessary to specify, and to use consistently, both a renormalization and a factorization scheme. The renormalization scheme used almost universally is the modified minimal subtraction ( $\overline{\text{MS}}$ ) scheme [41,42]. The most popular choices for the factorization scheme is also  $\overline{\text{MS}}$  [43]. However, sometimes the DIS [44] scheme is adopted, in which there are no higher-order corrections to the  $F_2$  structure function. The two schemes differ in how the non-divergent pieces are assimilated in the parton distribution functions.

The discussion above relates to the  $Q^2$  behavior of leading-twist (twist-2) contributions to the structure functions. Higher-twist terms, which involve their own non-perturbative input, exist. These die off as powers of  $Q$ ; specifically twist- $n$  terms are damped by  $1/Q^{n-2}$ . Provided a cut, say  $W^2 > 15 \text{ GeV}^2$  is imposed, the higher-twist terms appear to be numerically unimportant for  $Q^2$  above a few  $\text{GeV}^2$ , except for  $x$  close to 1 [45–47], though it is important to note that they are likely to be larger in  $xF_3(x, Q^2)$  than in  $F_2(x, Q^2)$  (see e.g. [48]).

### 19.3. Determination of parton distributions

The parton distribution functions (PDFs) can be determined from an analysis of data for deep inelastic lepton-nucleon scattering and for related hard-scattering processes initiated by nucleons; see [49–53] for reviews. Table 19.1 highlights some of the processes, where LHC data are playing an increasing role [54], and their primary sensitivity to PDFs. Fixed-target and collider experiments have complementary kinematic reach (as is shown in Fig. 19.4), which enables the determination of PDFs over a wide range in  $x$  and  $Q^2$ . As more precise LHC data for  $W^\pm, Z, \gamma, \text{jet}, b\bar{b}, t\bar{t}$  and  $J/\psi$  production become available, tighter constraints on the PDFs are expected in a wider kinematic range.

Recent determinations and releases of the unpolarized PDFs up to NNLO have been made by six groups: MMHT [55], NNPDF [56], CT(EQ) [57], HERAPDF [14], ABM [58] and JR [59]. JR generate ‘dynamical’ PDFs from a valence-like input at a very low



**Figure 19.4:** Kinematic domains in  $x$  and  $Q^2$  probed by fixed-target and collider experiments. Some of the final states accessible at the LHC are indicated in the appropriate regions, where  $y$  is the rapidity. The incoming partons have  $x_{1,2} = (M/14 \text{ TeV})e^{\pm y}$  with  $Q = M$  where  $M$  is the mass of the state shown in blue in the figure. For example, exclusive  $J/\psi$  and  $\Upsilon$  production at high  $|y|$  at the LHC may probe the gluon PDF down to  $x \sim 10^{-5}$ .

starting scale,  $Q_0^2 = 0.5 \text{ GeV}^2$ , whereas other groups start evolution at  $Q_0^2 = 1\text{--}4 \text{ GeV}^2$ . Most groups use input PDFs of the form  $xf = x^a(\dots)(1-x)^b$  with 14–28 free parameters in total. In these cases the PDF uncertainties are made available using the “Hessian” formulation. The free parameters are expanded around their best fit values, and orthogonal eigenvector sets of PDFs depending on linear combinations of the parameter variations are obtained. The uncertainty is then the quadratic sum of the uncertainties arising from each eigenvector. The NNPDF group combines a Monte Carlo representation of the probability measure in the space of PDFs with the use of neural networks. Fits are performed to a number of “replica” data sets obtained by allowing individual data points to fluctuate randomly by amounts determined by the size of the data uncertainties. This results in a set of replicas of unbiased PDF sets. In this case the best prediction is the average obtained using all PDF replicas and the uncertainty is the standard deviation over all replicas. It is now possible to convert the eigenvectors of Hessian-based PDFs to Monte Carlo replicas [60] and *vice versa* [61]. PDFs are made available in a common format at LHAPDF [62].

In these analyses the  $u, d$  and  $s$  quarks are taken to be massless, but the treatment of the heavy  $c$  and  $b$  quark masses,  $m_Q$ , differs, and has a long history, which may be traced from Refs. [63–74]. The MSTW, CT, NNPDF and HERAPDF analyses use different variants of the General-Mass Variable-Flavour-Number Scheme (GM-VFNS). This combines fixed-order contributions to the coefficient functions (or partonic cross sections) calculated with the full  $m_Q$  dependence, with the all-order resummation of contributions via DGLAP evolution in which the heavy quarks are treated as massless after starting evolution at some transition point. Transition matrix elements are computed, following [66], which provide the boundary conditions between  $n_f$  and  $n_f + 1$  PDFs. The ABM and JR analyses use a FFNS where only the three light (massless) quarks enter the evolution, while the heavy quarks enter the partonic cross sections with their full  $m_Q$  dependence. The GM-VFNS and FFNS approaches yield different results: in particular  $\alpha_s(M_Z^2)$  and the large- $x$  gluon PDF at large  $Q^2$  are both significantly smaller in the FFNS. It has been argued [46,47,73] that the difference is due to the slow convergence of the  $\ln^n(Q^2/m_Q^2)$  terms in certain regions in a FFNS.

The most recent determinations of the groups fitting a variety of data and using a GM-VFNS (MMHT, NNPDF and CT) have converged, so that now a good agreement has been achieved between the resulting PDFs. Indeed, the CT [57], MMHT [55], and NNPDF [56] PDF sets have been combined [75] using the Monte Carlo approach [60] mentioned above. The single combined set of PDFs is discussed in detail in Ref. [75].

For illustration, we show in Fig. 19.5 the PDFs obtained in the NNLO NNPDF analysis [56] at scales  $\mu^2 = 10$  and  $10^4 \text{ GeV}^2$ . The values of  $\alpha_s$  found by MMHT [76] may be taken as representative of those resulting from the GM-VFNS analyses

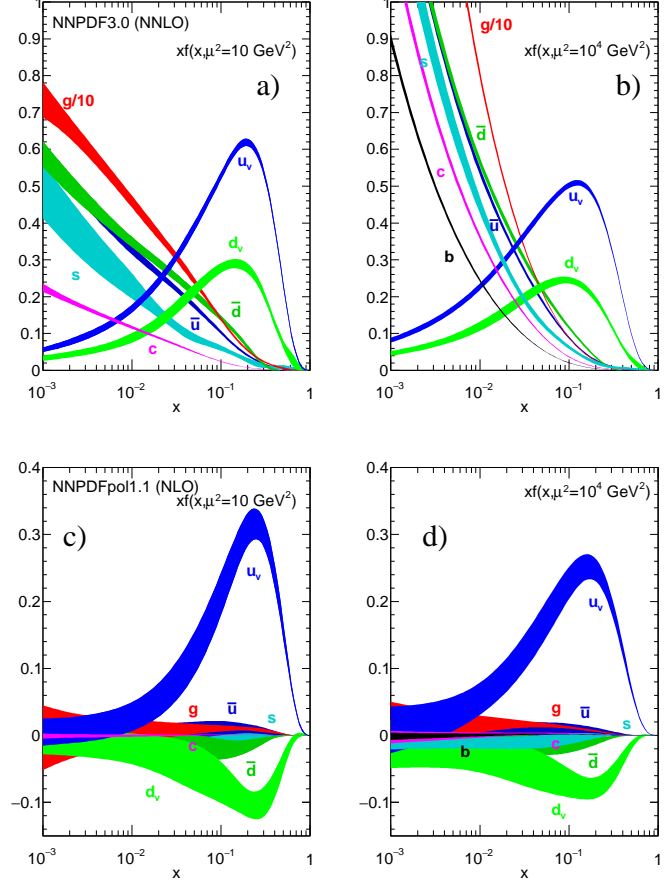
$$\text{NLO} : \alpha_s(M_Z^2) = 0.1201 \pm 0.0015,$$

$$\text{NNLO} : \alpha_s(M_Z^2) = 0.1172 \pm 0.0012,$$

where the error (at 68% C.L.) corresponds to the uncertainties resulting from the data fitted (the uncertainty that might be expected from the neglect of higher orders is at least as large), see also [77]. The ABM analysis [58], which uses a FFNS, finds  $\alpha_s(M_Z^2) = 0.1132 \pm 0.0011$  at NNLO.

Spin-dependent (or polarized) PDFs have been obtained through NLO global analyses which include measurements of the  $g_1$  structure function in inclusive polarized DIS, ‘flavour-tagged’ semi-inclusive DIS data, open-charm production in DIS and results from polarized  $pp$  scattering at RHIC. There are some very recent results on DIS from JLAB [78] and CLAS [79]. NLO analyses are given in Refs. [16–18] and [80,81]. Improved parton-to-hadron fragmentation functions, needed to describe the semi-inclusive DIS data, can be found in [82–84]. A recent determination [85], using the NNPDF methodology, concentrates just on the inclusive polarized DIS data, and finds the errors on the polarized gluon PDF have been underestimated in the earlier analyses. An update to this [15], where

jet and  $W^\pm$  data from  $pp$  collisions and open-charm DIS data have been included via reweighting reduces the uncertainty a little and suggests a positive polarized gluon PDF. The PDFs obtained in the NLO NNPDF analysis [15] at scales of  $\mu^2 = 10$  and  $10^4 \text{ GeV}^2$  are shown in Fig. 19.5.



**Figure 19.5:** The bands are  $x$  times the unpolarized (a,b) parton distributions  $f(x)$  (where  $f = u_v, d_v, \bar{u}, \bar{d}, s \simeq \bar{s}, c = \bar{c}, b = \bar{b}, g$ ) obtained in NNLO NNPDF3.0 global analysis [56] at scales  $\mu^2 = 10 \text{ GeV}^2$  (left) and  $\mu^2 = 10^4 \text{ GeV}^2$  (right), with  $\alpha_s(M_Z^2) = 0.118$ . The analogous results obtained in the NNLO MMHT analysis can be found in Fig. 1 of Ref [55]. The corresponding polarized parton distributions are shown (c,d), obtained in NLO with NNPDFpol1.1 [15].

Comprehensive sets of PDFs are available as program-callable functions from the HepData website [86], which includes comparison graphics of PDFs, and from the LHAPDF library [62], which can be linked directly into a user’s programme to provide access to recent PDFs in a standard format.

#### 19.4. The hadronic structure of the photon

Besides the *direct* interactions of the photon, it is possible for it to fluctuate into a hadronic state via the process  $\gamma \rightarrow q\bar{q}$ . While in this state, the partonic content of the photon may be *resolved*, for example, through the process  $e^+e^- \rightarrow e^+e^-\gamma^* \rightarrow e^+e^-X$ , where the virtual photon emitted by the DIS lepton probes the hadronic structure of the quasi-real photon emitted by the other lepton. The perturbative LO contributions,  $\gamma \rightarrow q\bar{q}$  followed by  $\gamma^*q \rightarrow g$ , are subject to QCD corrections due to the coupling of quarks to gluons.

Often the equivalent-photon approximation is used to express the differential cross section for deep inelastic electron-photon scattering in terms of the structure functions of the transverse quasi-real photon times a flux factor  $N_\gamma^T$  (for these incoming quasi-real photons of

transverse polarization)

$$\frac{d^2\sigma}{dx dQ^2} = N_\gamma^T \frac{2\pi\alpha^2}{xQ^4} \left[ (1 + (1-y)^2) F_2^\gamma(x, Q^2) - y^2 F_L^\gamma(x, Q^2) \right],$$

where we have used  $F_2^\gamma = 2xF_T^\gamma + F_L^\gamma$ , not to be confused with  $F_2^\gamma$  of Sec. 19.2. Complete formulae are given, for example, in the comprehensive review of Ref. 88.

The hadronic photon structure function,  $F_2^\gamma$ , evolves with increasing  $Q^2$  from the ‘hadron-like’ behavior, calculable via the vector-meson-dominance model, to the dominating ‘point-like’ behaviour, calculable in perturbative QCD. Due to the point-like coupling, the logarithmic evolution of  $F_2^\gamma$  with  $Q^2$  has a *positive* slope for all values of  $x$ , see Fig. 19.15. The ‘loss’ of quarks at large  $x$  due to gluon radiation is over-compensated by the ‘creation’ of quarks via the point-like  $\gamma \rightarrow q\bar{q}$  coupling. The logarithmic evolution was first predicted in the quark–parton model ( $\gamma^*\gamma \rightarrow q\bar{q}$ ) [89,90], and then in QCD in the limit of large  $Q^2$  [91]. The evolution is now known to NLO [92–94]. The NLO data analyses to determine the parton densities of the photon can be found in [95–97].

### 19.5. Diffractive DIS (DDIS)

Some 10% of DIS events are diffractive,  $\gamma^*p \rightarrow X + p$ , in which the slightly deflected proton and the cluster  $X$  of outgoing hadrons are well-separated in rapidity. Besides  $x$  and  $Q^2$ , two extra variables are needed to describe a DDIS event: the fraction  $x_{\mathbb{P}}$  of the proton’s momentum transferred across the rapidity gap and  $t$ , the square of the 4-momentum transfer of the proton. The DDIS data [98,99] are usually analyzed using two levels of factorization. First, the diffractive structure function  $F_2^D$  satisfies *collinear factorization*, and can be expressed as the convolution [100]

$$F_2^D = \sum_{a=q,g} C_2^a \otimes f_{a/p}^D, \quad (19.31)$$

with the same coefficient functions as in DIS (see Eq. (19.21)), and where the diffractive parton distributions  $f_{a/p}^D$  ( $a = q, g$ ) satisfy DGLAP evolution. Second, *Regge factorization* is assumed [101],

$$f_{a/p}^D(x_{\mathbb{P}}, t, z, \mu^2) = f_{\mathbb{P}/p}(x_{\mathbb{P}}, t) f_{a/\mathbb{P}}(z, \mu^2), \quad (19.32)$$

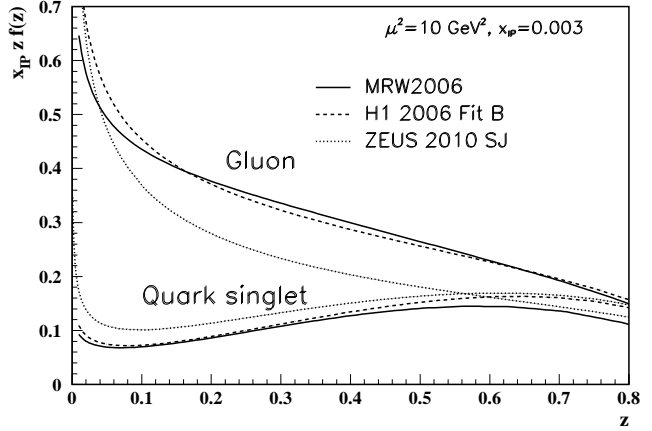
where  $f_{a/\mathbb{P}}$  are the parton densities of the Pomeron, which itself is treated like a hadron, and  $z \in [x/x_{\mathbb{P}}, 1]$  is the fraction of the Pomeron’s momentum carried by the parton entering the hard subprocess. The Pomeron flux factor  $f_{\mathbb{P}/p}(x_{\mathbb{P}}, t)$  is taken from Regge phenomenology. There are also secondary Reggeon contributions to Eq. (19.32). A sample of the  $t$ -integrated diffractive parton densities, obtained in this way, is shown in Fig. 19.6.

Although collinear factorization holds as  $\mu^2 \rightarrow \infty$ , there are non-negligible corrections for finite  $\mu^2$  and small  $x_{\mathbb{P}}$ . Besides the *resolved* interactions of the Pomeron, the perturbative QCD Pomeron may also interact *directly* with the hard subprocess, giving rise to an inhomogeneous evolution equation for the diffractive parton densities analogous to the photon case. The results of the MRW analysis [104], which includes these contributions, are also shown in Fig. 19.6. Unlike the inclusive case, the diffractive parton densities cannot be directly used to calculate diffractive hadron-hadron cross sections, since account must first be taken of ‘soft’ rescattering effects.

### 19.6. Generalized parton distributions

The parton distributions of the proton of Sec. 19.3 are given by the diagonal matrix elements  $\langle P, \lambda | \hat{O} | P, \lambda \rangle$ , where  $P$  and  $\lambda$  are the 4-momentum and helicity of the proton, and  $\hat{O}$  is a twist-2 quark or gluon operator. However, there is new information in the so-called generalised parton distributions (GPDs) defined in terms of the off-diagonal matrix elements  $\langle P', \lambda' | \hat{O} | P, \lambda \rangle$ ; see [106–110] for reviews. Unlike the diagonal PDFs, the GPDs cannot be regarded as parton densities, but are to be interpreted as probability amplitudes.

The physical significance of GPDs is best seen using light-cone coordinates,  $z^\pm = (z^0 \pm z^3)/\sqrt{2}$ , and in the light-cone gauge,  $A^+ = 0$ .



**Figure 19.6:** Diffractive parton distributions,  $x_{\mathbb{P}} z f_{a/p}^D$ , obtained from fitting to the ZEUS data with  $Q^2 > 5 \text{ GeV}^2$  [102], H1 data with  $Q^2 > 8.5 \text{ GeV}^2$  assuming Regge factorization [103], and from MRW2006 [104] using a more perturbative QCD approach [104]. Only the Pomeron contributions are shown and not the secondary Reggeon contributions, which are negligible at the value of  $x_{\mathbb{P}} = 0.003$  chosen here. The H1 2007 Jets distribution [105] is similar to H1 2006 Fit B.

It is conventional to define the generalised quark distributions in terms of quark operators at light-like separation

$$F_q(x, \xi, t) = \frac{1}{2} \int \frac{dz^-}{2\pi} e^{ix\bar{P}^+ z^-} \langle P' | \bar{\psi}(-z/2) \gamma^+ \psi(z/2) | P \rangle \Big|_{z^+ = z^1 = z^2 = 0} \quad (19.33)$$

$$= \frac{1}{2P^+} \left( H_q(x, \xi, t) \bar{u}(P') \gamma^+ u(P) + E_q(x, \xi, t) \bar{u}(P') \frac{i\sigma^{+\alpha} \Delta_\alpha}{2m} u(P) \right) \quad (19.34)$$

with  $\bar{P} = (P + P')/2$  and  $\Delta = P' - P$ , and where we have suppressed the helicity labels of the protons and spinors. We now have two extra kinematic variables:

$$t = \Delta^2, \quad \xi = -\Delta^+ / (P + P')^+. \quad (19.35)$$

We see that  $-1 \leq \xi \leq 1$ . Similarly, we may define GPDs  $\tilde{H}_q$  and  $\tilde{E}_q$  with an additional  $\gamma_5$  between the quark operators in Eq. (19.33); and also an analogous set of gluon GPDs,  $H_g$ ,  $E_g$ ,  $\tilde{H}_g$  and  $\tilde{E}_g$ . After a Fourier transform with respect to the transverse components of  $\Delta$ , we are able to describe the spatial distribution of partons in the impact parameter plane in terms of GPDs [111,112].

For  $P' = P$ ,  $\lambda' = \lambda$  the matrix elements reduce to the ordinary PDFs of Sec. 19.2.1

$$H_q(x, 0, 0) = q(x), \quad H_q(-x, 0, 0) = -\bar{q}(x), \quad H_g(x, 0, 0) = xg(x), \quad (19.36)$$

$$\tilde{H}_q(x, 0, 0) = \Delta q(x), \quad \tilde{H}_q(-x, 0, 0) = \Delta \bar{q}(x), \quad \tilde{H}_g(x, 0, 0) = x\Delta g(x), \quad (19.37)$$

where  $\Delta q = q \uparrow - q \downarrow$  as in Eq. (19.18). No corresponding relations exist for  $E$ ,  $\tilde{E}$  as they decouple in the forward limit,  $\Delta = 0$ .

The functions  $H_g, E_g$  are even in  $x$ , and  $\tilde{H}_g, \tilde{E}_g$  are odd functions of  $x$ . We can introduce valence and ‘singlet’ quark distributions which are even and odd functions of  $x$  respectively. For example

$$H_q^V(x, \xi, t) \equiv H_q(x, \xi, t) + H_q(-x, \xi, t) = H_q^V(-x, \xi, t), \quad (19.38)$$

$$H_q^S(x, \xi, t) \equiv H_q(x, \xi, t) - H_q(-x, \xi, t) = -H_q^S(-x, \xi, t). \quad (19.39)$$

All the GPDs satisfy relations of the form

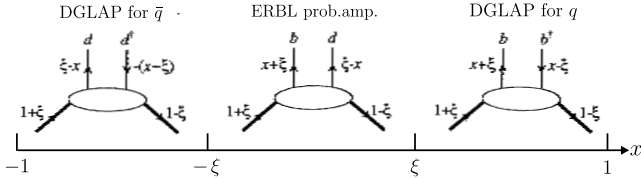
$$H(x, -\xi, t) = H(x, \xi, t) \quad \text{and} \quad H(x, -\xi, t)^* = H(x, \xi, t), \quad (19.40)$$



and so are real-valued functions. Moreover, the moments of GPDs, that is the  $x$  integrals of  $x^n H_q$  etc., are *polynomials* in  $\xi$  of order  $n+1$ . Another important property of GPDs are Ji's sum rules [106]

$$\frac{1}{2} \int_{-1}^1 dx \, x (H_q(x, \xi, t) + E_q(x, \xi, t)) = J_q(t), \quad (19.41)$$

where  $J_q(0)$  is the total angular momentum carried by quarks and antiquarks of flavour  $q$ , with a similar relation for gluons.



**Figure 19.7:** Schematic diagrams of the three distinct kinematic regions of the imaginary part of  $H_q$ . The proton and quark momentum fractions refer to  $\bar{P}^+$ , and  $x$  covers the interval  $(-1,1)$ . In the ERBL domain the GPDs are generalisations of distribution amplitudes which occur in processes such as  $p\bar{p} \rightarrow J/\psi$ .

To visualize the physical content of  $H_q$ , we Fourier expand  $\psi$  and  $\bar{\psi}$  in terms of quark, antiquark creation ( $b, d$ ) and annihilation ( $b^\dagger, d^\dagger$ ) operators, and sketch the result in Fig. 19.7. There are two types of domain: (i) the time-like or ‘annihilation’ domain, with  $|x| < |\xi|$ , where the GPDs describe the wave functions of a  $t$ -channel  $q\bar{q}$  (or gluon) pair and evolve according to modified ERBL equations [113,114]; (ii) the space-like or ‘scattering’ domain, with  $|x| > |\xi|$ , where the GPDs generalise the familiar  $\bar{q}, q$  (and gluon) PDFs and describe processes such as ‘deeply virtual Compton scattering’ ( $\gamma^* p \rightarrow \gamma p$ ),  $\gamma p \rightarrow J/\psi p$ , etc., and evolve according to modified DGLAP equations. The splitting functions for the evolution of GPDs are known to NLO [115].

GPDs describe new aspects of proton structure and must be determined from experiment. We can parametrise them in terms of ‘double distributions’ [116,117], which reduce to diagonal PDFs as  $\xi \rightarrow 0$ . With an additional physically reasonable ‘Regge’ assumption of no extra singularity at  $\xi = 0$ , GPDs at low  $\xi$  are uniquely given in terms of diagonal PDFs of  $O(\xi)$ , and have been used [118] to describe  $\gamma p \rightarrow J/\psi p$  data. Alternatively, flexible  $SO(3)$ -based parametrisations have been used to determine GPDs from DVCS data [119]; a more recent summary may be found in [120].

\* The value of  $\eta^{\text{CC}}$  deduced from Ref. 1 is found to be a factor of two too small;  $\eta^{\text{CC}}$  of Eq. (19.9) agrees with Refs. [2,3].

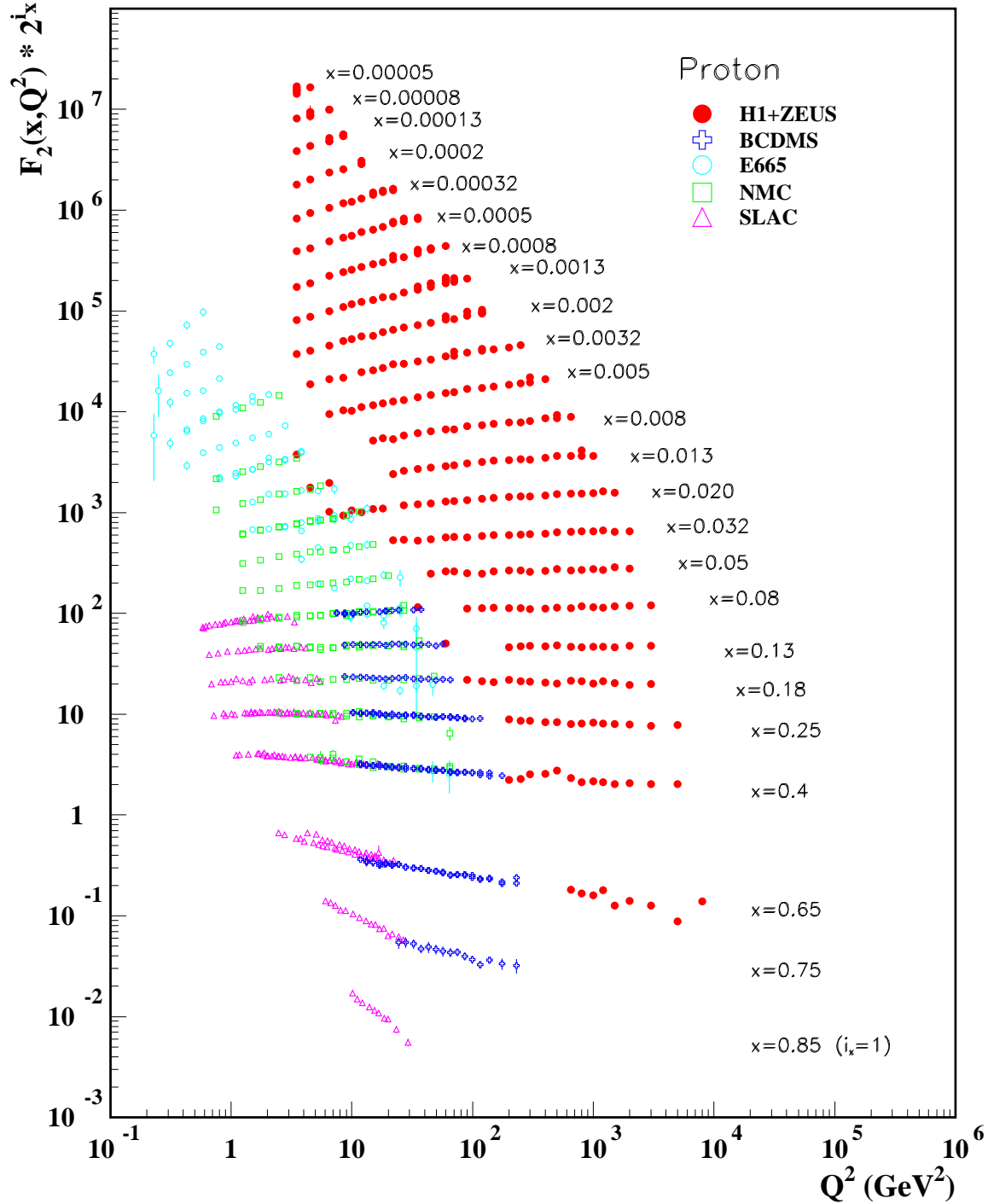
#### References:

1. J. Blümlein and N. Kochelev, Nucl. Phys. **B498**, 285 (1997).
2. S. Forte *et al.*, Nucl. Phys. **B602**, 585 (2001).
3. M. Anselmino *et al.*, Z. Phys. **C64**, 267 (1994).
4. M. Anselmino *et al.*, Phys. Rep. **261**, 1 (1995).
5. M. Klein and T. Riemann, Z. Phys. **C24**, 151 (1984).
6. C.G. Callan and D.J. Gross, Phys. Rev. Lett. **22**, 156 (1969).
7. D.A. Dicus, Phys. Rev. **D5**, 1367 (1972).
8. J.D. Bjorken and E.A. Paschos, Phys. Rev. **185**, 1975 (1969).
9. R.P. Feynman, Photon Hadron Interactions (Benjamin, New York, 1972).
10. S. Wandzura and F. Wilczek, Phys. Rev. **B72**, 195 (1977).
11. J. Blümlein and A. Tkabladze, Nucl. Phys. **B553**, 427 (1999).
12. J.D. Bjorken, Phys. Rev. **179**, 1547 (1969).
13. A.M. Cooper-Sarkar, private communication.
14. H1 and ZEUS Collaborations, H. Abramowicz *et al.*, Eur. Phys. J. **C75**, 580 (2015).
15. NNPDF, E.R. Nocera *et al.*, Nucl. Phys. **B887**, 276 (2014).
16. M. Hirai *et al.*, Nucl. Phys. **B813**, 106 (2009).
17. D. de Florian *et al.*, Phys. Rev. Lett. **101**, 072001 (2008); D. de Florian *et al.*, Phys. Rev. **D80**, 034030 (2009).
18. E. Leader *et al.*, Phys. Rev. **D82**, 114018 (2010).
19. HERMES, A. Airpetian *et al.*, Phys. Rev. Lett. **92**, 012005 (2004); A. Airpetian *et al.*, Phys. Rev. **D71**, 012003 (2005).
20. HERMES, A. Airpetian *et al.*, Phys. Lett. **B666**, 446 (2008).
21. SMC, B. Adeva *et al.*, Phys. Lett. **B420**, 180 (1998).
22. COMPASS, M. Alekseev *et al.*, Phys. Lett. **B680**, 217 (2009).
23. COMPASS, M. Alekseev *et al.*, Phys. Lett. **B693**, 227 (2010).
24. V.N. Gribov and L.N. Lipatov, Sov. J. Nucl. Phys. **15**, 438 (1972).
25. L.N. Lipatov, Sov. J. Nucl. Phys. **20**, 95 (1975).
26. G. Altarelli and G. Parisi, Nucl. Phys. **B126**, 298 (1977).
27. Yu.L. Dokshitzer, Sov. Phys. JETP **46**, 641 (1977).
28. G. Curci *et al.*, Nucl. Phys. **B175**, 27 (1980); W. Furmanski, and R. Petronzio, Phys. Lett. **B97**, 437 (1980).
29. R.K. Ellis *et al.*, QCD and Collider Physics (Cambridge UP, 1996).
30. E.B. Zijlstra and W.L. van Neerven, Phys. Lett. **B272**, 127 (1991); E.B. Zijlstra and W.L. van Neerven, Phys. Lett. **B273**, 476 (1991); E.B. Zijlstra and W.L. van Neerven, Phys. Lett. **B297**, 377 (1992); E.B. Zijlstra and W.L. van Neerven, Nucl. Phys. **B383**, 525 (1992).
31. S. Moch and J.A.M. Vermaseren, Nucl. Phys. **B573**, 853 (2000).
32. S. Moch *et al.*, Nucl. Phys. **B688**, 101 (2004); S. Moch *et al.*, Nucl. Phys. **B691**, 129 (2004); S. Moch *et al.*, Phys. Lett. **B606**, 123 (2005); S. Moch *et al.*, Nucl. Phys. **B724**, 3 (2005).
33. E.A. Kuraev *et al.*, Phys. Lett. **B60**, 50 (1975); E.A. Kuraev *et al.*, Sov. Phys. JETP **44**, 443 (1976); E.A. Kuraev *et al.*, Sov. Phys. JETP **45**, 199 (1977).
34. Ya.Ya. Balitsky and L.N. Lipatov, Sov. J. Nucl. Phys. **28**, 822 (1978).
35. V.S. Fadin, and L.N. Lipatov, Phys. Lett. **B429**, 127 (1998).
36. G. Camici and M. Ciafaloni, Phys. Lett. **B412**, 396 (1997), erratum-Phys. Lett. **B147**, 390 (1997); G. Camici and M. Ciafaloni, Phys. Lett. **B430**, 349 (1998).
37. M. Ciafaloni *et al.*, Phys. Rev. **D60**, 114036 (1999); M. Ciafaloni *et al.*, JHEP **0007**, 054 (2000).
38. M. Ciafaloni *et al.*, Phys. Lett. **B576**, 143 (2003); M. Ciafaloni *et al.*, Phys. Rev. **D68**, 114003 (2003).
39. G. Altarelli *et al.*, Nucl. Phys. **B742**, 1 (2006); G. Altarelli *et al.*, Nucl. Phys. **B799**, 199 (2008).
40. C.D. White and R.S. Thorne, Phys. Rev. **D75**, 034005 (2007).
41. G. 't Hooft and M. Veltman, Nucl. Phys. **B44**, 189 (1972).
42. G. 't Hooft, Nucl. Phys. **B61**, 455 (1973).
43. W.A. Bardeen *et al.*, Phys. Rev. **D18**, 3998 (1978).
44. G. Altarelli *et al.*, Nucl. Phys. **B143**, 521 (1978) and erratum: Nucl. Phys. **B146**, 544 (1978).
45. A.D. Martin *et al.*, Eur. Phys. J. **C35**, 325 (2004).
46. NNPDF, R.D. Ball *et al.*, Phys. Lett. **B723**, 330 (2013).
47. R.S. Thorne, Eur. Phys. J. **C74**, 2958 (2014).
48. M. Dasgupta and B.R. Webber, Phys. Lett. **B382**, 273 (1996).
49. A. De Roeck and R.S. Thorne, Prog. in Part. Nucl. Phys. **66**, 727 (2011).
50. S. Forte and G. Watt, Ann. Rev. Nucl. and Part. Sci. **63**, 291 (2013).
51. J. Blumlein, Prog. in Part. Nucl. Phys. **69**, 28 (2013).
52. E. Perez and E. Rizvi, Rept. on Prog. in Phys. **76**, 046201 (2013).
53. R.D. Ball *et al.*, JHEP **1304**, 125 (2013).
54. J. Rojo *et al.*, J. Phys. **G42**, 103103 (2015).
55. MMHT, L. Harland-Lang *et al.*, Eur. Phys. J. **C75**, 204 (2015).
56. NNPDF, R.D. Ball *et al.*, JHEP **1504**, 40 (2015).
57. CT14, S. Dulat *et al.*, arXiv:1506.07443.
58. S. Alekhin, *et al.*, Phys. Rev. **D89**, 054028 (2014).

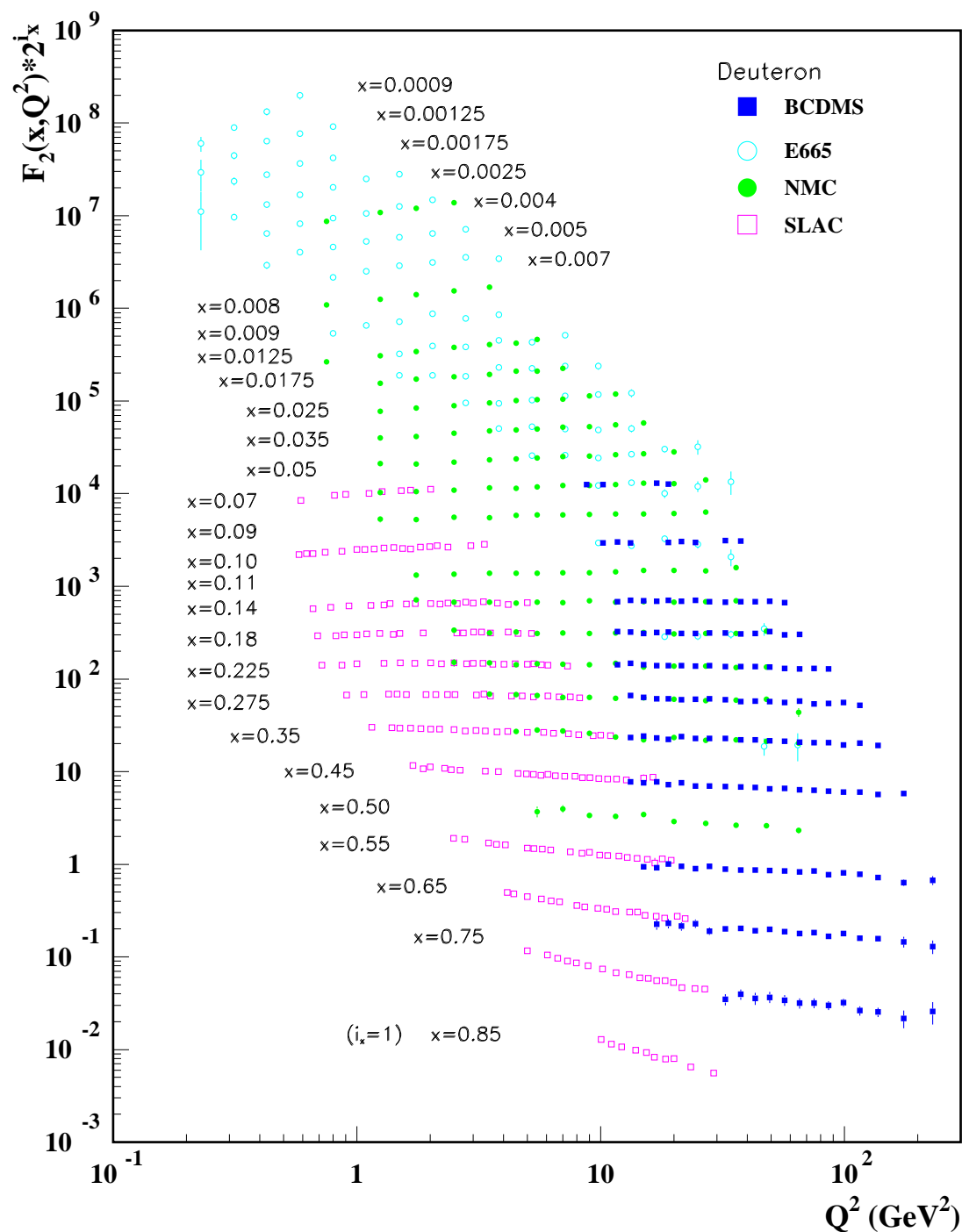
59. P. Jimenez-Delgado and E. Reya, Phys. Rev. **D89**, 074049 (2014).
60. G. Watt and R.S. Thorne, JHEP **1208**, 052 (2012).
61. S. Carrazza *et al.*, Eur. Phys. J. **C75**, 369 (2015).
62. A. Buckley *et al.*, Eur. Phys. J. **C75**, 132 (2015).
63. J.C. Collins, F. Wilczek, and A. Zee, Phys. Rev. **D18**, 242 (1978).
64. E. Laenen *et al.*, Nucl. Phys. **B392**, 162 (1993).
65. M.A.G. Aivazis *et al.*, Phys. Rev. **D50**, 3102 (1994).
66. M. Buza *et al.*, Eur. Phys. J. **C1**, 301 (1998).
67. J.C. Collins, Phys. Rev. **D58**, 094002 (1998).
68. A. Chuvakin *et al.*, Phys. Rev. **D61**, 096004 (2000).
69. R.S. Thorne, Phys. Rev. **D73**, 054019 (2006).
70. R.S. Thorne and W.-K. Tung, Proc. 4th HERA-LHC Workshop, [arXiv:0809.0714](https://arxiv.org/abs/0809.0714).
71. S. Alekhin and S. Moch, Phys. Lett. **B699**, 345 (2011).
72. S. Forte *et al.*, Nucl. Phys. **B834**, 116 (2010).
73. R.S. Thorne, Phys. Rev. **D86**, 074017 (2012).
74. E.G. de Oliveira *et al.*, Eur. Phys. J. **C73**, 2616 (2013).
75. J. Butterworth *et al.*, J. Phys. **G43**, 023001 (2016).
76. MMHT, L. Harland-Lang *et al.*, Eur. Phys. J. **C75**, 435 (2015).
77. NNPDF, R.D. Ball *et al.*, Phys. Lett. **B707**, 66 (2012).
78. Jefferson Lab Hall A Collaboration, D.S. Parno *et al.*, Phys. Lett. **B744**, 309 (2015).
79. CLAS, Y. Prok *et al.*, Phys. Rev. **C90**, 025212 (2014).
80. J. Blümlein and H. Böttcher, Nucl. Phys. **B841**, 205 (2010).
81. P. Jimenez-Delgado *et al.*, Phys. Rev. **D89**, 034025 (2014).
82. D. de Florian *et al.*, Phys. Rev. **D75**, 114010 (2007);  
D. de Florian *et al.*, Phys. Rev. **D76**, 074033 (2007).
83. S. Albino *et al.*, Nucl. Phys. **B803**, 42 (2008).
84. M. Hirai and S. Kumano, Comp. Phys. Comm. **183**, 1002 (2012).
85. NNPDF, R.D. Ball *et al.*, Nucl. Phys. **B874**, 36 (2013).
86. <http://hepdata.cedar.ac.uk/pdfs>.
87. <http://lhpdf.hepforge.org/>.
88. R. Nisius, Phys. Reports **332**, 165 (2000).
89. T.F. Walsh and P.M. Zerwas, Phys. Lett. **B44**, 195 (1973).
90. R.L. Kingsley, Nucl. Phys. **B60**, 45 (1973).
91. E. Witten, Nucl. Phys. **B120**, 189 (1977).
92. W.A. Bardeen and A.J. Buras, Phys. Rev. **D20**, 166 (1979),  
erratum Phys. Rev. **D21**, 2041 (1980).
93. M. Fontannaz and E. Pilon, Phys. Rev. **D45**, 382 (1992),  
erratum Phys. Rev. **D46**, 484 (1992).
94. M. Glück *et al.*, Phys. Rev. **D45**, 3986 (1992).
95. F. Cornet *et al.*, Phys. Rev. **D70**, 093004 (2004).
96. P. Aurenche, *et al.*, Eur. Phys. J. **C44**, 395 (2005).
97. W. Slominski *et al.*, Eur. Phys. J. **C45**, 633 (2006).
98. H1+ZEUS, F.D. Aaron *et al.*, Eur. Phys. J. **C72**, 2175 (2012).
99. H1, F.D. Aaron *et al.*, Eur. Phys. J. **C72**, 2074 (2012).
100. J.C. Collins, Phys. Rev. **D57**, 3051 (1998); erratum Phys. Rev. **D61**, 019902 (2000).
101. G. Ingelman and P. E. Schlein, Phys. Lett. **B152**, 256 (1985).
102. ZEUS, S. Chekanov *et al.*, Nucl. Phys. **B831**, 1 (2010).
103. H1, A. Aktas *et al.*, Eur. Phys. J. **C48**, 715 (2006).
104. A.D. Martin, M.G. Ryskin, and G. Watt, Phys. Lett. **B644**, 131 (2007).
105. H1, A. Aktas *et al.*, JHEP **0710**, 042 (2007).
106. X. Ji, J. Phys. **G24**, 1181 (1998).
107. K. Goeke *et al.*, Prog. in Part. Nucl. Phys. **47**, 401 (2001).
108. M. Diehl, Phys. Rept. **388**, 41 (2003).
109. A.V. Belitsky and A.V. Radyushkin, Phys. Rept. **418**, 1 (2005).
110. S. Boffi and B. Pasquini, Riv. Nuovo Cimento **30**, 387 (2007).
111. M. Burkardt, Int. J. Mod. Phys. **A18**, 173 (2003).
112. M. Diehl, Eur. Phys. J. **C25**, 223 (2002).
113. A.V. Efremov and A.V. Radyushkin, Phys. Lett. **B94**, 245 (1980).
114. G.P. Lepage and S.J. Brodsky, Phys. Rev. **D22**, 2157 (1980).
115. A.V. Belitsky *et al.*, Phys. Lett. **B493**, 341 (2000).
116. A.V. Radyushkin, Phys. Rev. **D59**, 014030 (1999).
117. A.V. Radyushkin, Phys. Lett. **B449**, 81 (1999).
118. A.D. Martin *et al.*, Eur. Phys. J. **C63**, 57 (2009).
119. K. Kumericki and D. Müller, Nucl. Phys. **B841**, 1 (2010).
120. M. Guidal *et al.*, Rept. Prog. Phys. **76** 066202 (2013).



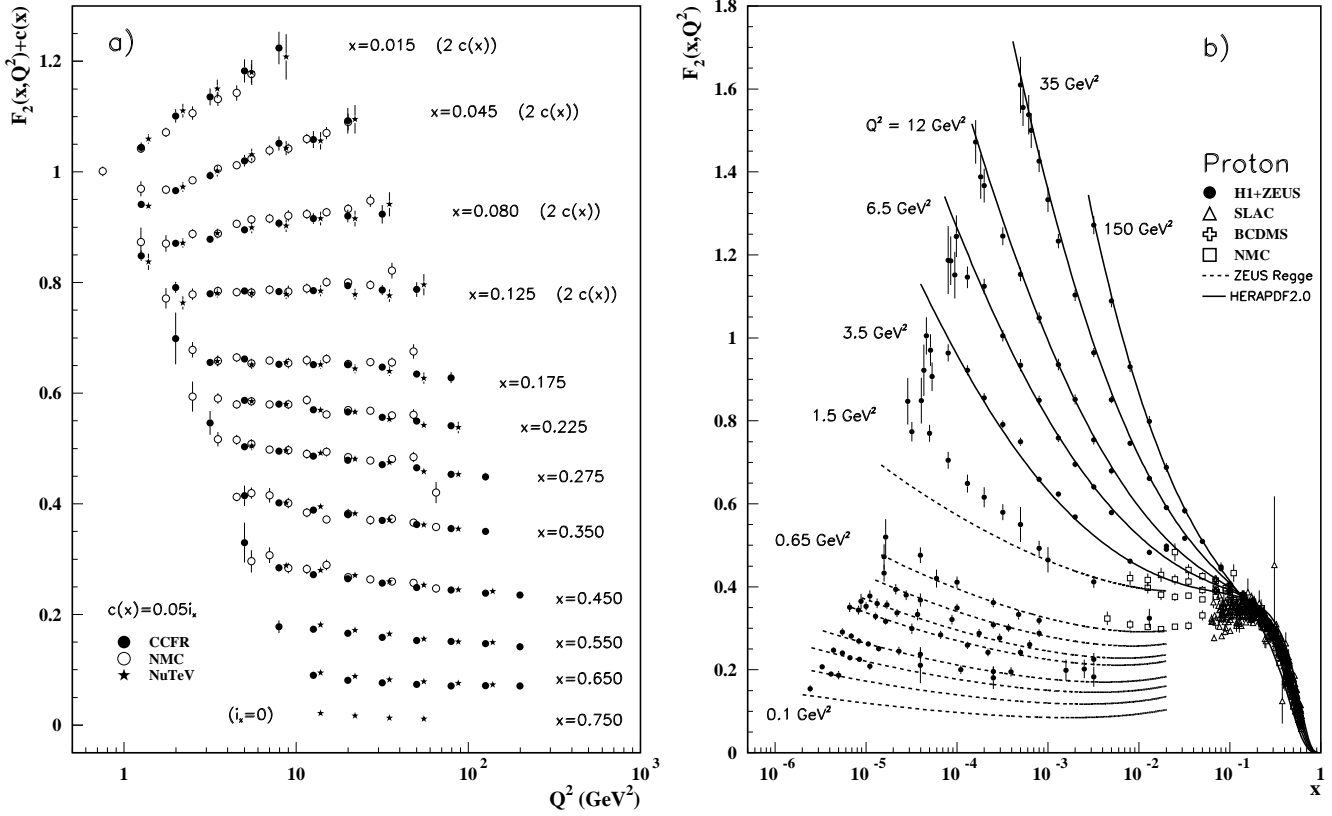
NOTE: THE FIGURES IN THIS SECTION ARE INTENDED TO SHOW THE REPRESENTATIVE DATA. THEY ARE NOT MEANT TO BE COMPLETE COMPILATIONS OF ALL THE WORLD'S RELIABLE DATA.



**Figure 19.8:** The proton structure function  $F_2^p$  measured in electromagnetic scattering of electrons and positrons on protons (collider experiments H1 and ZEUS for  $Q^2 \geq 2 \text{ GeV}^2$ ), in the kinematic domain of the HERA data (see Fig. 19.10 for data at smaller  $x$  and  $Q^2$ ), and for electrons (SLAC) and muons (BCDMS, E665, NMC) on a fixed target. Statistical and systematic errors added in quadrature are shown. The H1+ZEUS combined values are obtained from the measured reduced cross section and converted to  $F_2^p$  with a HERAPDF NLO fit, for all measured points where the predicted ratio of  $F_2^p$  to reduced cross-section was within 10% of unity. The data are plotted as a function of  $Q^2$  in bins of fixed  $x$ . Some points have been slightly offset in  $Q^2$  for clarity. The H1+ZEUS combined binning in  $x$  is used in this plot; all other data are rebinned to the  $x$  values of these data. For the purpose of plotting,  $F_2^p$  has been multiplied by  $2^{i_x}$ , where  $i_x$  is the number of the  $x$  bin, ranging from  $i_x = 1$  ( $x = 0.85$ ) to  $i_x = 24$  ( $x = 0.00005$ ). References: **H1 and ZEUS**—H. Abramowicz *et al.*, Eur. Phys. J. **C75**, 580 (2015) (for both data and HERAPDF parameterization); **BCDMS**—A.C. Benvenuti *et al.*, Phys. Lett. **B223**, 485 (1989) (as given in [86]); **E665**—M.R. Adams *et al.*, Phys. Rev. **D54**, 3006 (1996); **NMC**—M. Arneodo *et al.*, Nucl. Phys. **B483**, 3 (1997); **SLAC**—L.W. Whitlow *et al.*, Phys. Lett. **B282**, 475 (1992).



**Figure 19.9:** The deuteron structure function  $F_2^d$  measured in electromagnetic scattering of electrons (SLAC) and muons (BCDMS, E665, NMC) on a fixed target, shown as a function of  $Q^2$  for bins of fixed  $x$ . Statistical and systematic errors added in quadrature are shown. For the purpose of plotting,  $F_2^d$  has been multiplied by  $2^{i_x}$ , where  $i_x$  is the number of the  $x$  bin, ranging from 1 ( $x = 0.85$ ) to 29 ( $x = 0.0009$ ). References: **BCDMS**—A.C. Benvenuti *et al.*, Phys. Lett. **B237**, 592 (1990). **E665**, **NMC**, **SLAC**—same references as Fig. 19.8.

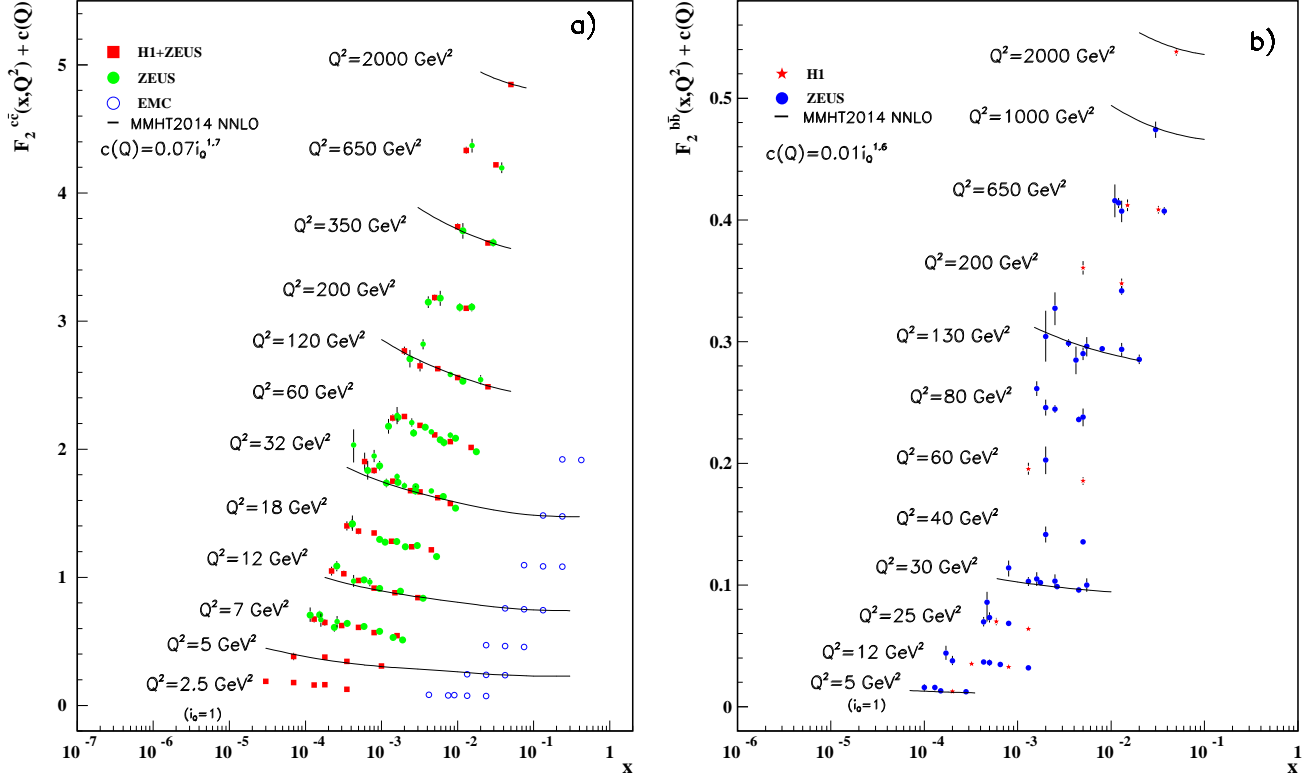


**Figure 19.10:** a) The deuteron structure function  $F_2$  measured in deep inelastic scattering of muons on a fixed target (NMC) is compared to the structure function  $F_2$  from neutrino-iron scattering (CCFR and NuTeV) using  $F_2^\mu = (5/18)F_2^p - x(s + \bar{s})/6$ , where heavy-target effects have been taken into account. The data are shown versus  $Q^2$ , for bins of fixed  $x$ . The NMC data have been rebinned to CCFR and NuTeV  $x$  values. For the purpose of plotting, a constant  $c(x) = 0.05i_x$  is added to  $F_2$ , where  $i_x$  is the number of the  $x$  bin, ranging from 0 ( $x = 0.75$ ) to 7 ( $x = 0.175$ ). For  $i_x = 8$  ( $x = 0.125$ ) to 11 ( $x = 0.015$ ),  $2c(x)$  has been added. References: **NMC**—M. Arneodo *et al.*, Nucl. Phys. **B483**, 3 (1997); **CCFR/NuTeV**—U.K. Yang *et al.*, Phys. Rev. Lett. **86**, 2741 (2001); **NuTeV**—M. Tzanov *et al.*, Phys. Rev. **D74**, 012008 (2006).

b) The proton structure function  $F_2^p$  mostly at small  $x$  and  $Q^2$ , measured in electromagnetic scattering of electrons and positrons (H1, ZEUS), electrons (SLAC), and muons (BCDMS, NMC) on protons. Lines are ZEUS Regge and HERAPDF parameterizations for lower and higher  $Q^2$ , respectively. The width of the bins can be up to 10% of the stated  $Q^2$ . Some points have been slightly offset in  $x$  for clarity. The H1+ZEUS combined values for  $Q^2 \geq 3.5$  GeV<sup>2</sup> are obtained from the measured reduced cross section and converted to  $F_2^p$  with a HERAPDF NLO fit, for all measured points where the predicted ratio of  $F_2^p$  to reduced cross-section was within 10% of unity. A turn-over is visible in the low- $x$  points at medium  $Q^2$  (3.5 GeV<sup>2</sup> and 6 GeV<sup>2</sup>) for the H1+ZEUS combined values. In order to obtain  $F_2^p$  from the measured reduced cross-section,  $F_L$  must be estimated; for the points shown, this estimate is obtained from HERAPDF2.0. No  $F_L$  value consistent with the HERA data can eliminate the turn-over. This may indicate that at low  $x$  and  $Q^2$  there are contributions to the structure functions that cannot be described in standard DGLAP evolution.

References: **H1 and ZEUS**—F.D. Aaron *et al.*, JHEP **1001**, 109 (2010) (data for  $Q^2 < 3.5$  GeV<sup>2</sup>), H. Abramowicz *et al.*, Eur. Phys. J. **C75**, 580 (2015) (data for  $Q^2 \geq 3.5$  GeV<sup>2</sup> and HERAPDF parameterization); **ZEUS**—J. Breitweg *et al.*, Phys. Lett. **B487**, 53 (2000) (ZEUS Regge parameterization); **BCDMS, NMC, SLAC**—same references as Fig. 19.8.

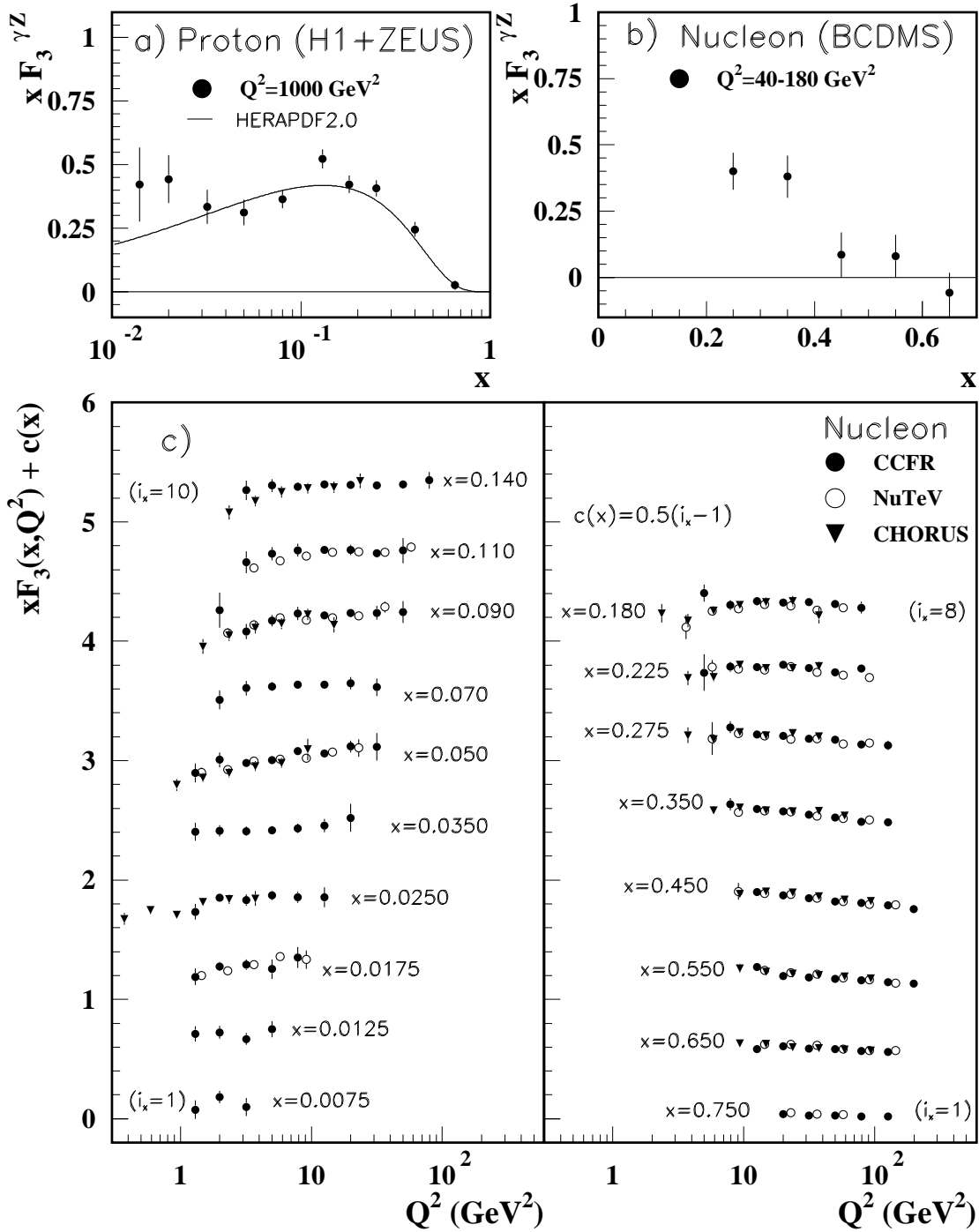
Statistical and systematic errors added in quadrature are shown for both plots.



**Figure 19.11:** a) The charm-quark structure function  $F_2^{c\bar{c}}(x)$ , i.e. that part of the inclusive structure function  $F_2^p$  arising from the production of charm quarks, measured in electromagnetic scattering of positrons on protons (H1, ZEUS) and muons on iron (EMC). For the purpose of plotting, a constant  $c(Q) = 0.07i_Q^{1.7}$  is added to  $F_2^{c\bar{c}}$  where  $i_Q$  is the number of the  $Q^2$  bin, ranging from 1 ( $Q^2 = 2.5 \text{ GeV}^2$ ) to 12 ( $Q^2 = 2000 \text{ GeV}^2$ ). References: **H1 and ZEUS run I combination**—H. Abramowicz *et al.*, Eur. Phys. J. **C73**, 2311 (2013); **ZEUS run II**—H. Abramowicz *et al.*, JHEP **05**, 023 (2013); H. Abramowicz *et al.*, JHEP **05**, 097 (2013); H. Abramowicz *et al.*, JHEP **09**, 127 (2014); **EMC**—J.J. Aubert *et al.*, Nucl. Phys. **B213**, 31 (1983).

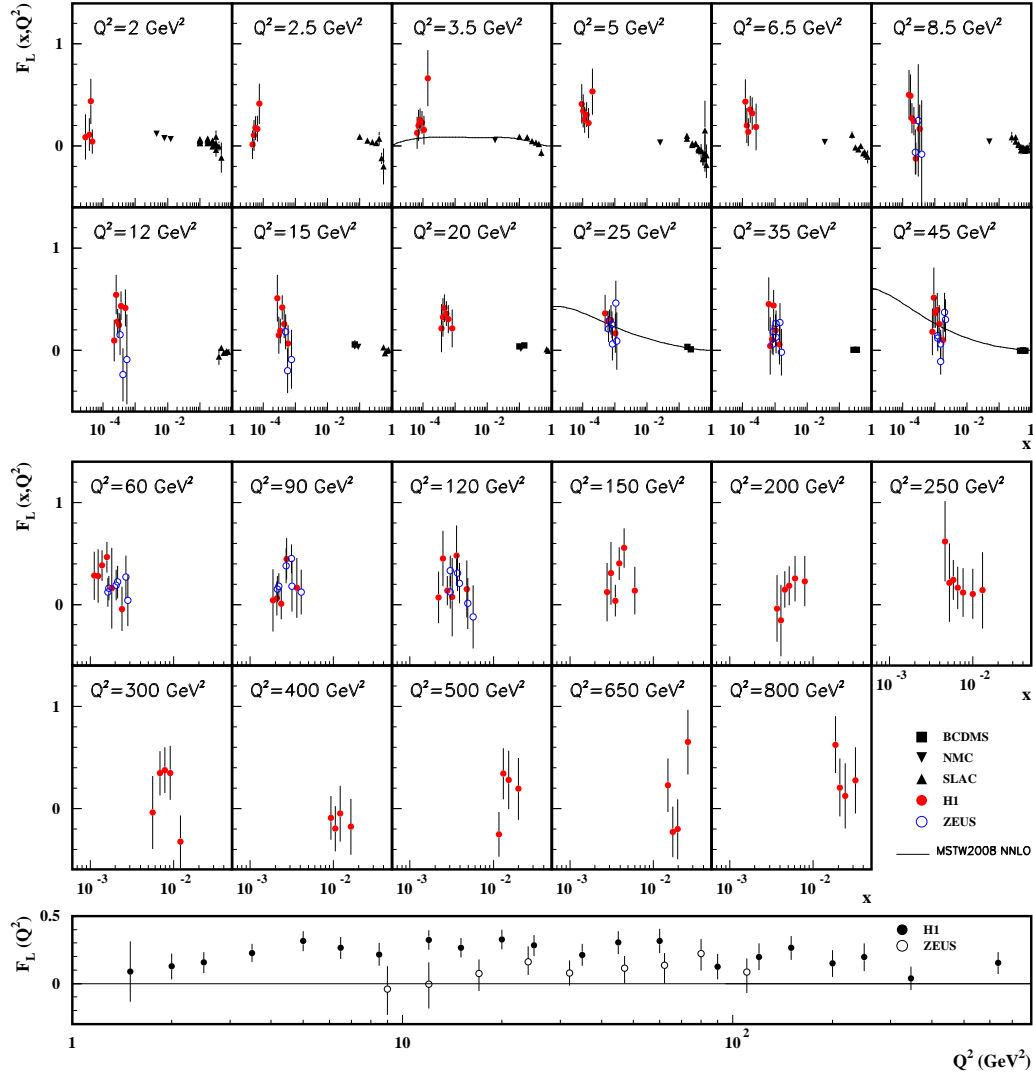
b) The bottom-quark structure function  $F_2^{b\bar{b}}(x)$ . For the purpose of plotting, a constant  $c(Q) = 0.01i_Q^{1.6}$  is added to  $F_2^{b\bar{b}}$  where  $i_Q$  is the number of the  $Q^2$  bin, ranging from 1 ( $Q^2 = 5 \text{ GeV}^2$ ) to 12 ( $Q^2 = 2000 \text{ GeV}^2$ ). References: **ZEUS**—S. Chekanov *et al.*, Eur. Phys. J. **C65**, 65 (2010); H. Abramowicz *et al.*, Eur. Phys. J. **C69**, 347 (2010); H. Abramowicz *et al.*, Eur. Phys. J. **C71**, 1573 (2011); H. Abramowicz *et al.*, JHEP **09**, 127 (2014); **H1**—F.D. Aaron *et al.*, Eur. Phys. J. **C65**, 89 (2010).

For both plots, statistical and systematic errors added in quadrature are shown. The data are given as a function of  $x$  in bins of  $Q^2$ . Points may have been slightly offset in  $x$  for clarity. Some data have been rebinned to common  $Q^2$  values. Also shown is the MMHT2014 parameterization given at several  $Q^2$  values (L. A. Harland-Lang *et al.*, Eur. Phys. J. **C75**, 204 (2015)).



**Figure 19.12:** The structure function  $x F_3^{\gamma Z}$  measured in electroweak scattering of **a)** electrons on protons (H1 and ZEUS) and **b)** muons on carbon (BCDMS). The line in **a)** is the HERAPDF parameterization. References: **H1 and ZEUS**—H. Abramowicz *et al.*, Eur. Phys. J. **C75**, 580 (2015) (for both data and HERAPDF parameterization); **BCDMS**—A. Argento *et al.*, Phys. Lett. **B140**, 142 (1984). **c)** The structure function  $x F_3$  of the nucleon measured in  $\nu$ -Fe scattering. The data are plotted as a function of  $Q^2$  in bins of fixed  $x$ . For the purpose of plotting, a constant  $c(x) = 0.5(i_x - 1)$  is added to  $x F_3$ , where  $i_x$  is the number of the  $x$  bin as shown in the plot. The NuTeV and CHORUS points have been shifted to the nearest corresponding  $x$  bin as given in the plot and slightly offset in  $Q^2$  for clarity. References: **CCFR**—W.G. Seligman *et al.*, Phys. Rev. Lett. **79**, 1213 (1997); **NuTeV**—M. Tzanov *et al.*, Phys. Rev. **D74**, 012008 (2006); **CHORUS**—G. Önençüt *et al.*, Phys. Lett. **B632**, 65 (2006).

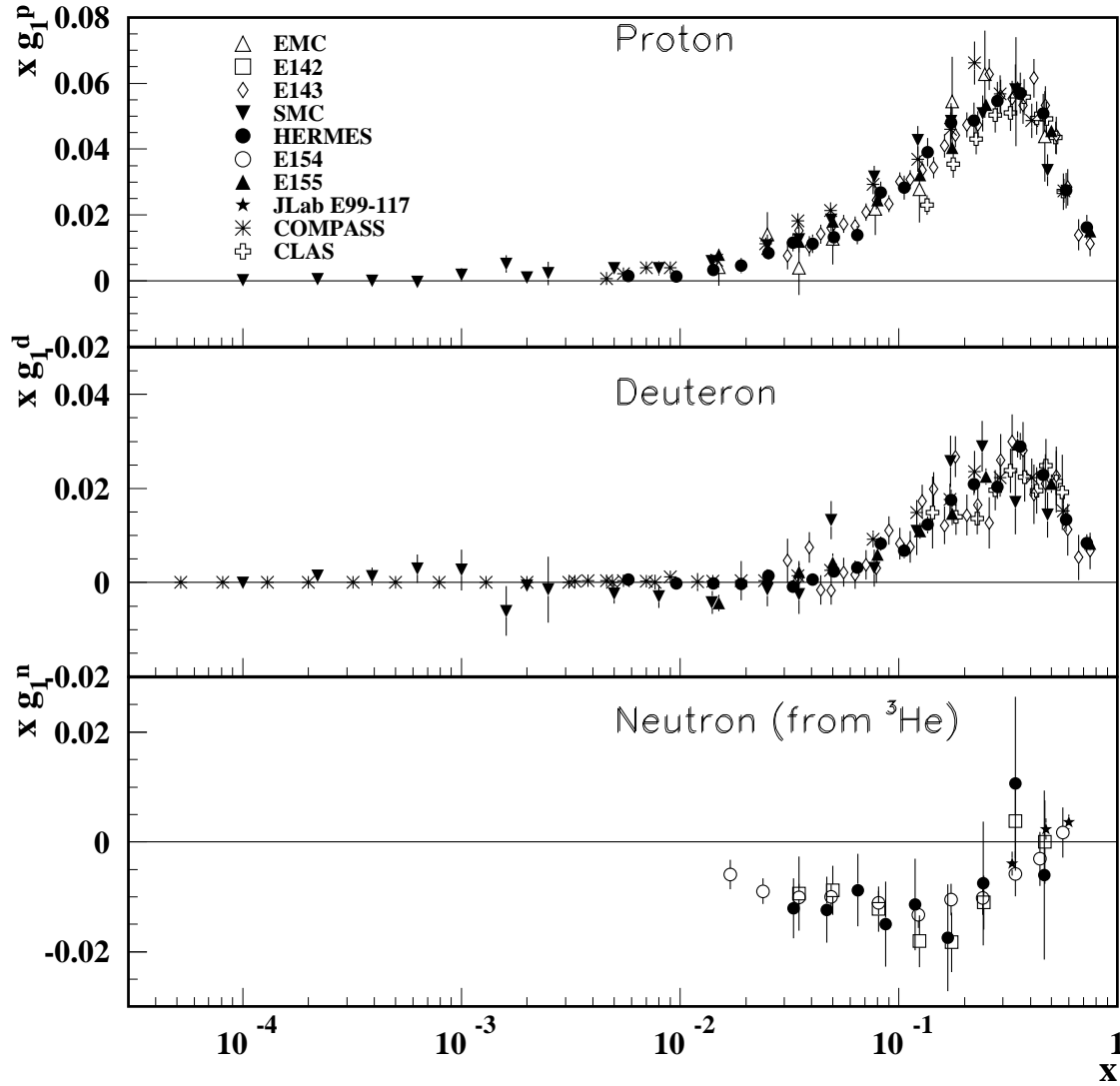
Statistical and systematic errors added in quadrature are shown for all plots.



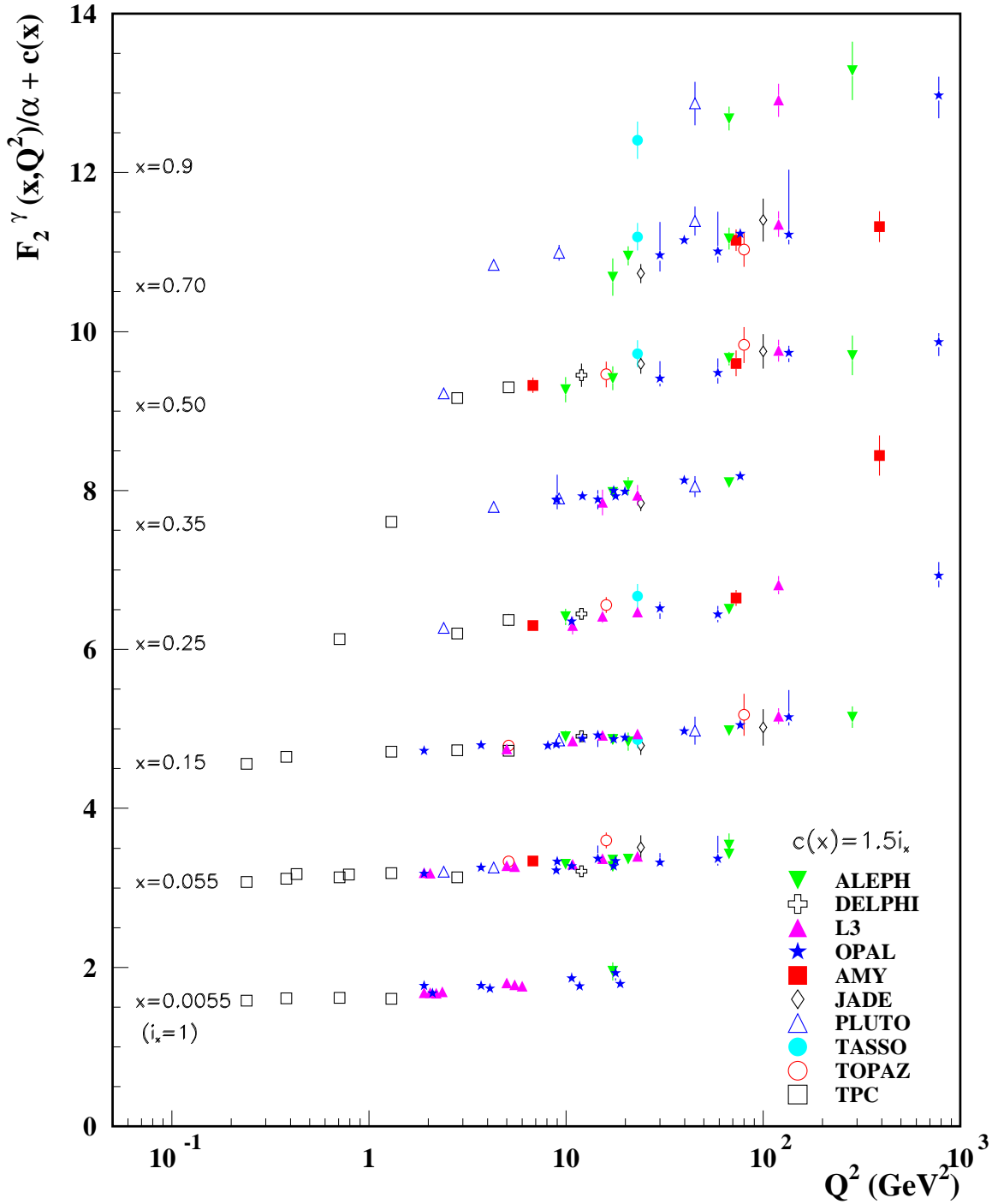
**Figure 19.13:** Top panels: The longitudinal structure function  $F_L$  as a function of  $x$  in bins of fixed  $Q^2$  measured on the proton (except for the SLAC data which also contain deuterium data). BCDMS, NMC, and SLAC results are from measurements of  $R$  (the ratio of longitudinal to transverse photon absorption cross sections) which are converted to  $F_L$  by using the BDCMS parameterization of  $F_2$  (A.C. Benvenuti *et al.*, Phys. Lett. **B223**, 485 (1989)). It is assumed that the  $Q^2$  dependence of the fixed-target data is small within a given  $Q^2$  bin. Some of the other data may have been rebinned to common  $Q^2$  values. Some points have been slightly offset in  $x$  for clarity. Also shown is the MSTW2008 parameterization given at three  $Q^2$  values (A.D. Martin *et al.*, Eur. Phys. J. **C63**, 189 (2009)). References: **H1**—V. Andreev *et al.*, Eur. Phys. J. **C74**, 2814 (2014); **ZEUS**—S. Chekanov *et al.*, Phys. Lett. **B682**, 8 (2009); H. Abramowicz *et al.*, Phys. Rev. **D90**, 072002 (2014); **BCDMS**—A. Benvenuti *et al.*, Phys. Lett. **B223**, 485 (1989); **NMC**—M. Arneodo *et al.*, Nucl. Phys. **B483**, 3 (1997); **SLAC**—L.W. Whitlow *et al.*, Phys. Lett. **B250**, 193 (1990) and numerical values from the thesis of L.W. Whitlow (SLAC-357).

Bottom panel: The longitudinal structure function  $F_L$  as a function of  $Q^2$ . Some points have been slightly offset in  $Q^2$  for clarity. References: **H1**—V. Andreev *et al.*, Eur. Phys. J. **C74**, 2814 (2014); **ZEUS**—H. Abramowicz *et al.*, Phys. Rev. **D90**, 072002 (2014).

The results shown in the bottom plot require the assumption of the validity of the QCD form for the  $F_2$  structure function in order to extract  $F_L$ . Statistical and systematic errors added in quadrature are shown for both plots.



**Figure 19.14:** The spin-dependent structure function  $xg_1(x)$  of the proton, deuteron, and neutron (from  $^3\text{He}$  target) measured in deep inelastic scattering of polarized electrons/positrons: E142 ( $Q^2 \sim 0.3 - 10 \text{ GeV}^2$ ), E143 ( $Q^2 \sim 0.3 - 10 \text{ GeV}^2$ ), E154 ( $Q^2 \sim 1 - 17 \text{ GeV}^2$ ), E155 ( $Q^2 \sim 1 - 40 \text{ GeV}^2$ ), JLab E99-117 ( $Q^2 \sim 2.71 - 4.83 \text{ GeV}^2$ ), HERMES ( $Q^2 \sim 0.18 - 20 \text{ GeV}^2$ ), CLAS ( $Q^2 \sim 1 - 5 \text{ GeV}^2$ ) and muons: EMC ( $Q^2 \sim 1.5 - 100 \text{ GeV}^2$ ), SMC ( $Q^2 \sim 0.01 - 100 \text{ GeV}^2$ ), COMPASS ( $Q^2 \sim 0.001 - 100 \text{ GeV}^2$ ), shown at the measured  $Q^2$  (except for EMC data given at  $Q^2 = 10.7 \text{ GeV}^2$  and E155 data given at  $Q^2 = 5 \text{ GeV}^2$ ). Note that  $g_1^n(x)$  may also be extracted by taking the difference between  $g_1^d(x)$  and  $g_1^p(x)$ , but these values have been omitted in the bottom plot for clarity. Statistical and systematic errors added in quadrature are shown. References: **EMC**—J. Ashman *et al.*, Nucl. Phys. **B328**, 1 (1989); **E142**—P.L. Anthony *et al.*, Phys. Rev. **D54**, 6620 (1996); **E143**—K. Abe *et al.*, Phys. Rev. **D58**, 112003 (1998); **SMC**—B. Adeva *et al.*, Phys. Rev. **D58**, 112001 (1998), B. Adeva *et al.*, Phys. Rev. **D60**, 072004 (1999) and Erratum-Phys. Rev. **D62**, 079902 (2000); **HERMES**—A. Airapetian *et al.*, Phys. Rev. **D75**, 012007 (2007) and K. Ackerstaff *et al.*, Phys. Lett. **B404**, 383 (1997); **E154**—K. Abe *et al.*, Phys. Rev. Lett. **79**, 26 (1997); **E155**—P.L. Anthony *et al.*, Phys. Lett. **B463**, 339 (1999) and P.L. Anthony *et al.*, Phys. Lett. **B493**, 19 (2000); **Jlab-E99-117**—X. Zheng *et al.*, Phys. Rev. **C70**, 065207 (2004); **COMPASS**—V.Yu. Alexakhin *et al.*, Phys. Lett. **B647**, 8 (2007), E.S. Ageev *et al.*, Phys. Lett. **B647**, 330 (2007), and M.G. Alekseev *et al.*, Phys. Lett. **B690**, 466 (2010); **CLAS**—K.V. Dharmawardane *et al.*, Phys. Lett. **B641**, 11 (2006) (which also includes resonance region data not shown on this plot).



**Figure 19.15:** The hadronic structure function of the photon  $F_2^\gamma$  divided by the fine structure constant  $\alpha$  measured in  $e^+e^-$  scattering, shown as a function of  $Q^2$  for bins of  $x$ . Data points have been shifted to the nearest corresponding  $x$  bin as given in the plot. Some points have been offset in  $Q^2$  for clarity. Statistical and systematic errors added in quadrature are shown. For the purpose of plotting, a constant  $c(x) = 1.5i_x$  is added to  $F_2^\gamma/\alpha$  where  $i_x$  is the number of the  $x$  bin, ranging from 1 ( $x = 0.0055$ ) to 8 ( $x = 0.9$ ). References: **ALEPH**—R. Barate *et al.*, Phys. Lett. **B458**, 152 (1999); A. Heister *et al.*, Eur. Phys. J. **C30**, 145 (2003); **DELPHI**—P. Abreu *et al.*, Z. Phys. **C69**, 223 (1995); **L3**—M. Acciarri *et al.*, Phys. Lett. **B436**, 403 (1998); M. Acciarri *et al.*, Phys. Lett. **B447**, 147 (1999); M. Acciarri *et al.*, Phys. Lett. **B483**, 373 (2000); **OPAL**—A. Ackerstaff *et al.*, Phys. Lett. **B411**, 387 (1997); A. Ackerstaff *et al.*, Z. Phys. **C74**, 33 (1997); G. Abbiendi *et al.*, Eur. Phys. J. **C18**, 15 (2000); G. Abbiendi *et al.*, Phys. Lett. **B533**, 207 (2002) (note that there is overlap of the data samples in these last two papers); **AMY**—S.K. Sahu *et al.*, Phys. Lett. **B346**, 208 (1995); T. Kojima *et al.*, Phys. Lett. **B400**, 395 (1997); **JADE**—W. Bartel *et al.*, Z. Phys. **C24**, 231 (1984); **PLUTO**—C. Berger *et al.*, Phys. Lett. **142B**, 111 (1984); C. Berger *et al.*, Nucl. Phys. **B281**, 365 (1987); **TASSO**—M. Althoff *et al.*, Z. Phys. **C31**, 527 (1986); **TOPAZ**—K. Muramatsu *et al.*, Phys. Lett. **B332**, 477 (1994); **TPC/Two Gamma**—H. Aihara *et al.*, Z. Phys. **C34**, 1 (1987).



## 20. FRAGMENTATION FUNCTIONS IN $e^+e^-$ , $ep$ AND $pp$ COLLISIONS

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### 20.1. Introduction to fragmentation

The term ‘fragmentation functions’ is widely used for two conceptually different (albeit related) sets of functions describing final-state single particle energy distributions in hard scattering processes (see Refs. [1,2] for introductory reviews, and Refs. [3,4] for summaries of experimental and theoretical research in this field).

The first are cross-section observables such as the functions  $F_{T,L,A}(x, s)$  in semi-inclusive  $e^+e^-$  annihilation at center-of-mass (CM) energy  $\sqrt{s}$  via an intermediate photon or  $Z$ -boson,  $e^+e^- \rightarrow \gamma/Z \rightarrow h + X$ , given by

$$\frac{1}{\sigma_0} \frac{d^2\sigma^h}{dx d\cos\theta} = \frac{3}{8}(1 + \cos^2\theta)F_T^h(x, s) + \frac{3}{4}\sin^2\theta F_L^h(x, s) + \frac{3}{4}\cos\theta F_A^h(x, s). \quad (20.1)$$

Here  $x = 2E_h/\sqrt{s} \leq 1$  is the scaled energy of the hadron  $h$  (in practice the approximation  $x \simeq x_p = 2p_h/\sqrt{s}$  or  $x \simeq p/p_{max}$  is often used), and  $\theta$  is its angle relative to the electron beam in the CM frame. Eq. (20.1) is the most general form for unpolarized inclusive single-particle production via vector bosons [5]. The transverse and longitudinal fragmentation functions  $F_T$  and  $F_L$  represent the contributions from  $\gamma/Z$  polarizations transverse or longitudinal with respect to the direction of motion of the hadron. The parity-violating term with the asymmetric fragmentation function  $F_A$  arises from the interference between vector and axial-vector contributions. Normalization factors  $\sigma_0$  used in the literature range from the total cross section  $\sigma_{tot}$  for  $e^+e^- \rightarrow$  hadrons, including all weak and QCD contributions, to  $\sigma_0 = 4\pi\alpha^2 N_c/3s$  with  $N_c = 3$ , the lowest-order QED cross section for  $e^+e^- \rightarrow \mu^+\mu^-$  times the number of colors  $N_c$ . LEP1 measurements of all three fragmentation functions are shown in Fig. 20.1.

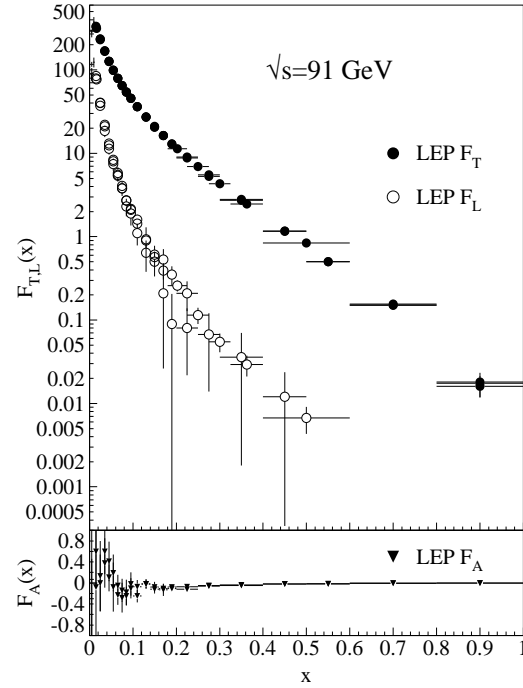
Integration of Eq. (20.1) over  $\theta$  yields the total fragmentation function  $F^h = F_T^h + F_L^h$ ,

$$\frac{1}{\sigma_0} \frac{d\sigma^h}{dx} = F^h(x, s) = \sum_i \int_x^1 \frac{dz}{z} C_i(z, \alpha_s(\mu), \frac{s}{\mu^2}) D_i^h(\frac{x}{z}, \mu^2) + \mathcal{O}(\frac{1}{\sqrt{s}}) \quad (20.2)$$

with  $i = u, \bar{u}, d, \bar{d}, \dots, g$ . Here the second set of functions mentioned in the first paragraph has been introduced, the parton fragmentation functions (or fragmentation densities)  $D_i^h$ . These functions are the final-state analogue of the initial-state parton distribution functions (pdf) addressed in Section 19 of this *Review*. Due to the different sign of the squared four-momentum  $q^2$  of the intermediate gauge boson these two sets of fragmentation distributions are also referred to as the timelike ( $e^+e^-$  annihilation,  $q^2 > 0$ ) and spacelike (deep-inelastic scattering (DIS),  $q^2 < 0$ ) parton distribution functions. The function  $D_i^h(z, \mu^2)$  describes the probability that the parton  $i$  fragments into a hadron  $h$  carrying a probability that the parton  $i$  fragments into a hadron  $h$  carrying a fraction  $z$  of the parton’s momentum. Beyond the leading order (LO) of perturbative QCD these universal functions are factorization-scheme dependent, with ‘reasonable’ scheme choices retaining certain quark-parton-model [6] (QPM) constraints such as the momentum sum rule

$$\sum_h \int_0^1 dz z D_i^h(z, \mu^2) = 1. \quad (20.3)$$

The dependence of the functions  $D_i^h$  on the factorization scale  $\mu^2$  is discussed in Section 20.2. Like in Eq. (20.2) and below, this scale is often taken to be equal to the factorization or renormalization scale, but this equivalence is not required in the theory.



**Figure 20.1:** LEP1 measurements of total transverse ( $F_T$ ), longitudinal ( $F_L$ ), and asymmetric ( $F_A$ ) fragmentation functions [7–9]. Data points with relative errors greater than 100% are omitted.

The second ingredient in Eq. (20.2), and analogous expressions for the functions  $F_{T,L,A}$ , are the observable-dependent coefficient functions  $C_i$ . At the zeroth order in the strong coupling  $\alpha_s$  the coefficient functions  $C_g$  for gluons are zero, while for (anti-)quarks  $C_i = g_i(s)\delta(1-z)$  except for  $F_L$ , where  $g_i(s)$  is the appropriate electroweak coupling. In particular,  $g_i(s)$  is proportional to the squared charge of the quark  $i$  at  $s \ll M_Z^2$ , when weak effects can be neglected. The full electroweak prefactors  $g_i(s)$  can be found in Ref. [5]. The power corrections in Eq. (20.2) arise from quark and hadron mass terms and from non-perturbative effects.

Measurements of fragmentation in lepton-hadron and hadron-hadron scattering are complementary to those in  $e^+e^-$  annihilation. The former are affected by contributions, in summary called the hadron remnant, arising from the partons of the initial-state hadron which are collaterally involved in the hard lepton-parton or parton-parton collision. The latter provides a clean environment (no initial-state hadron remnant) and stringent constraints on the combinations  $D_{q_i}^h + D_{\bar{q}_i}^h$ . However  $e^+e^-$  annihilation is far less sensitive to  $D_g^h$  and insensitive to the charge asymmetries  $D_{q_i}^h - D_{\bar{q}_i}^h$ . These quantities are best constrained in proton-(anti-)proton and electron-proton scattering, respectively. Especially the latter provides a more complicated environment with which it is possible to study the influence on the fragmentation process from initial-state QCD radiation, the partonic and spin structure of the hadron target, and the target remnant system (see Ref. [10] for a comprehensive review of the measurements and models of fragmentation in lepton-hadron scattering).

Moreover, unlike  $e^+e^-$  annihilation where  $q^2 = s$  is fixed by the collider energy, lepton-hadron scattering has two independent scales,  $Q^2 = -q^2$  and the invariant mass  $W^2$  of the hadronic final state, which both can vary by several orders of magnitudes for a given CM energy, thus allowing the study of fragmentation in different environments by a single experiment. E.g., in photoproduction the exchanged photon is quasi-real ( $Q^2 \approx 0$ ) leading to processes akin to hadron-hadron scattering. In DIS ( $Q^2 \gg 1 \text{ GeV}^2$ ), using the QPM, the hadronic fragments of the struck quark can be directly compared

with quark fragmentation in  $e^+e^-$  in a suitable frame. Results from lepton-hadron experiments quoted in this report primarily concern fragmentation in the DIS regime. Studies performed by lepton-hadron experiments of fragmentation with photoproduction data containing high transverse momentum jets or particles are also reported, when these are directly comparable to DIS and  $e^+e^-$  results.

Fragmentation studies in lepton-hadron collisions are usually performed in one of two frames in which the target hadron and the exchanged boson are collinear. The hadronic center-of-mass frame (HCMS) is defined as the rest system of the exchanged boson and incoming hadron, with the  $z^*$ -axis defined along the direction of the exchanged boson. The positive  $z^*$  direction defines the so-called current region. Fragmentation measurements performed in the HCMS often use the Feynman- $x$  variable  $x_F = 2p_z^*/W$ , where  $p_z^*$  is the longitudinal momentum of the particle in this frame. As  $W$  is the invariant mass of the hadronic final state,  $x_F$  ranges between  $-1$  and  $1$ .

The Breit system [11] is connected to the HCMS by a longitudinal boost such that the time component of  $q$  vanishes, i.e.  $q = (0, 0, 0, -Q)$ . In the QPM, the struck parton then has the longitudinal momentum  $Q/2$  which becomes  $-Q/2$  after the collision. As compared with the HCMS, the current region of the Breit frame is more closely matched to the partonic scattering process, and is thus appropriate for direct comparisons of fragmentation functions in DIS with those from  $e^+e^-$  annihilation. The variable  $x_p = 2p^*/Q$  is used at HERA for measurements in the Breit frame, ensuring rather directly comparable DIS and  $e^+e^-$  results, where  $p^*$  is the particle's momentum in the current region of the Breit frame.

## 20.2. Scaling violation

The simplest parton-model approach would predict scale-independent  $x$ -distributions ('scaling') for both the fragmentation function  $F^h$  and the parton fragmentation functions  $D_i^h$ . Perturbative QCD corrections lead, after factorization of the final-state collinear singularities for light partons, to logarithmic scaling violations via the evolution equations [12]

$$\frac{\partial}{\partial \ln \mu^2} D_i(x, \mu^2) = \sum_j \int_x^1 \frac{dz}{z} P_{ji}(z, \alpha_s(\mu^2)) D_j\left(\frac{x}{z}, \mu^2\right), \quad (20.4)$$

where the splitting functions  $P_{ij}(z, \alpha_s(\mu^2))$  describe in leading order the probability to find parton  $i$  with a longitudinal momentum fraction  $z$  in parton  $j$ . Usually this system of equations is decomposed into a  $2 \times 2$  flavour-singlet sector comprising gluon and the sum of all quark and antiquark fragmentation functions, and scalar ('non-singlet') equations for quark-antiquark and flavour differences. The singlet splitting-function matrix is now  $P_{ji}$ , rather than  $P_{ij}$  as for the initial-state parton distributions, since  $D_j$  represents the fragmentation of the final parton.

The splitting functions in Eq. (20.4) have perturbative expansion of the form

$$P_{ji}(z, \alpha_s) = \frac{\alpha_s}{2\pi} P_{ji}^{(0)}(z) + \left(\frac{\alpha_s}{2\pi}\right)^2 P_{ji}^{(1)}(z) + \left(\frac{\alpha_s}{2\pi}\right)^3 P_{ji}^{(2)}(z) + \dots \quad (20.5)$$

where the leading-order (LO) functions  $P^{(0)}(z)$  [12,13] are the same as those for the initial-state parton distributions. The next-to-leading order (NLO) corrections  $P^{(1)}(z)$  have been calculated in Refs. [14–18] (there are well-known misprints in the journal version of Ref. [15]). Ref. [18] also includes the spin-dependent case. These functions are different from, but related to their space-like counterparts, see also Ref. [19]. These relations have facilitated recent calculations of the next-to-next-to-leading order (NNLO) quantities  $P_{qq}^{(2)}(z)$  and  $P_{gg}^{(2)}(z)$  in Eq. (20.5) [20,21]. The corresponding off-diagonal quantities  $P_{qg}^{(2)}$  and  $P_{gq}^{(2)}$  were recently obtained in Ref. [22] by using similar relations supplemented with constraints from the momentum sum rule Eq. (20.3) [21] and from the limit of  $C_A = C_F = n_f$  for which QCD becomes supersymmetric. An uncertainty, which does not affect the logarithmic behaviour at small and large momentum fractions, still remains on the  $P_{qg}^{(2)}$  kernel. All these results refer to the standard  $\overline{\text{MS}}$

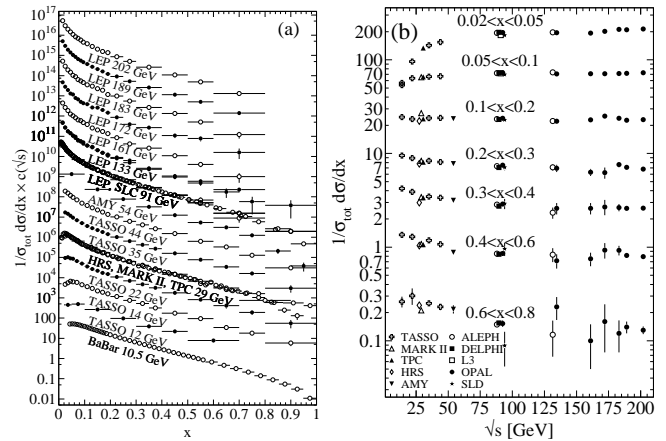
scheme, with the exception of Refs. [17], with a fixed number  $n_f$  of light flavours. Fragmentation functions change when in the course of energy evolution the threshold for the production of a heavier quark flavour is crossed. The NLO treatment of these flavour thresholds in the evolution has been addressed in Ref. [23].

The QCD parts of the coefficient functions for  $F_{T,L,A}(x, s)$  in Eq. (20.1) and the total fragmentation function  $F_2^h \equiv F^h$  in Eq. (20.2) are given by

$$C_{a,i}(z, \alpha_s) = (1 - \delta_{aL}) \delta_{iq} + \frac{\alpha_s}{2\pi} c_{a,i}^{(1)}(z) + \left(\frac{\alpha_s}{2\pi}\right)^2 c_{a,i}^{(2)}(z) + \dots \quad (20.6)$$

The first-order corrections have been calculated in Refs. [24], and the second-order terms in Ref. 25. The latter results have been verified (and some typos corrected) in Refs. [20,26]. The coefficient functions are known to NNLO except for  $F_L$  where the leading contribution is of order  $\alpha_s$ .

The effect of the evolution is similar in the timelike and spacelike cases: as the scale increases, one observes a scaling violation in which the  $x$ -distribution is shifted towards lower values. This can be seen from Fig. 20.2 where a large amount of measurements of the total fragmentation function in  $e^+e^-$  annihilation are summarized. QCD analyses of these data are discussed in Section 20.5 below.



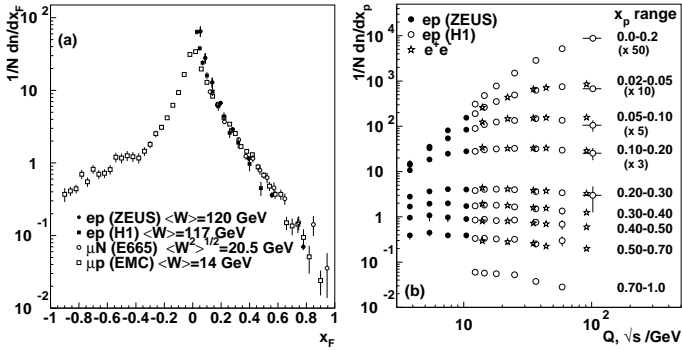
**Figure 20.2:** The  $e^+e^-$  fragmentation function for all charged particles is shown [9,27–44] (a) for different CM energies  $\sqrt{s}$  versus  $x$  and (b) for various ranges of  $x$  versus  $\sqrt{s}$ . For the purpose of plotting (a), the distributions were scaled by  $c(\sqrt{s}) = 10^i$  with  $i$  ranging from  $i = 0$  ( $\sqrt{s} = 12$  GeV) to  $i = 13$  ( $\sqrt{s} = 202$  GeV).

Unlike the splitting functions in Eq. (20.5), see Refs. [19–21], the coefficient functions for  $F_{2,T,A}$  in Eq. (20.6) show a threshold enhancement with terms up to  $\alpha_s^n (1-z)^{-1} \ln^{2n-1}(1-z)$ . Such logarithms can be resummed to all orders in  $\alpha_s$  using standard soft-gluon techniques [45–47]. Recently this resummation has been extended to the subleading (and for  $F_L$  leading) class  $\alpha_s^n \ln^k(1-z)$  of large- $x$  logarithms [48,49].

In Refs. [24] the NLO coefficient functions have been calculated also for single hadron production in lepton-proton scattering,  $ep \rightarrow e + h + X$ . More recently corresponding results have been obtained for the case that a non-vanishing transverse momentum is required in the HCMS frame [50].

Scaling violations in DIS are shown in Fig. 20.3 for both HCMS and Breit frame. In Fig. 1.3(a) the distribution in terms of  $x_F = 2p_z^*/W$  shows a steeper slope in  $ep$  data than for the lower-energy  $\mu p$  data for  $x_F > 0.15$ , indicating the scaling violations. At smaller values of  $x_F$  in the current jet region, the multiplicity of particles substantially increases with  $W$  owing to the increased phase space available for the fragmentation process. The EMC data access both the current region and the region of the fragmenting target remnant system. At higher values of  $|x_F|$ , due to the extended nature of the remnant, the

multiplicity in the target region far exceeds that in the current region. For acceptance reasons the remnant hemisphere of the HCMS is only accessible by the lower-energy fixed-target experiments.



**Figure 20.3:** (a) The distribution  $1/N \cdot dN/dx_F$  for all charged particles in DIS lepton-hadron experiments at different values of  $W$ , and measured in the HCMS [51–54]. (b) Scaling violations of the fragmentation function for all charged particles in the current region of the Breit frame of DIS [55,60] and in  $e^+e^-$  interactions [37,61]. The data are shown as a function of  $\sqrt{s}$  for  $e^+e^-$  results, and as a function of  $Q$  for the DIS results, each within the same indicated intervals of the scaled momentum  $x_p$ . The data for the four lowest intervals of  $x_p$  are multiplied by factors 50, 10, 5, and 3, respectively for clarity.

Using hadrons from the current hemisphere in the Breit frame, measurements of fragmentation functions and the production properties of particles in  $ep$  scattering have been made by Refs. [55–60]. Fig. 20.3(b) compares results from  $ep$  scattering and  $e^+e^-$  experiments, the latter results are halved as they cover both event hemispheres. The agreement between the DIS and  $e^+e^-$  results is fairly good. However, processes in DIS which are not present in  $e^+e^-$  annihilation, such as boson-gluon fusion and initial-state QCD radiation, can depopulate the current region. These effects become most prominent at low values of  $Q$  and  $x_p$ . Hence, when compared with  $e^+e^-$  annihilation data at  $\sqrt{s} = 5.2, 6.5$  GeV [62] not shown here, the DIS particle rates tend to lie below those from  $e^+e^-$  annihilation. A ZEUS study [63] finds that the direct comparability of the  $ep$  data to  $e^+e^-$  results at low scales is improved if twice the energy in the current hemisphere of the Breit frame,  $2E_B^{CT}$ , is used instead of  $Q/2$  as the fragmentation scale. Choosing  $2 \cdot E_B^{CT}$  for the fragmentation scale approximates QCD radiation effects relevant at low scales as detailed in Ref. [64].

### 20.3. Fragmentation functions for small particle momenta

The higher-order timelike splitting functions in Eq. (20.5) are very singular at small  $x$ . They show a double-logarithmic (LL) enhancement with leading terms of the form  $\alpha_s^n \ln^{2n-2} x$  corresponding to poles  $\alpha_s^n (N-1)^{1-2n}$  for the Mellin moments

$$P^{(n)}(N) = \int_0^1 dx x^{N-1} P^{(n)}(x). \quad (20.7)$$

Despite large cancellations between leading and non-leading logarithms at non-asymptotic value of  $x$ , the resulting small- $x$  rise in the timelike splitting functions dwarfs that of their spacelike counterparts for the evolution of the parton distributions in Section 19 of this Review, see Fig. 1 of Ref. [21]. Consequently the fixed-order approximation to the evolution breaks down orders of magnitude in  $x$  earlier in fragmentation than in DIS.

The pattern of the known coefficients and other considerations suggest that the LL terms sum to all-order expressions without any pole at  $N = 1$  such as [65,66]

$$P_{gg}^{LL}(N) = -\frac{1}{4}(N-1 - \sqrt{(N-1)^2 + 24\alpha_s/\pi}). \quad (20.8)$$

Keeping the first three terms in the resulting expansion of Eq. (20.4) around  $N = 1$  yields a Gaussian in the variable  $\xi = \ln(1/x)$  for the small- $x$  fragmentation functions,

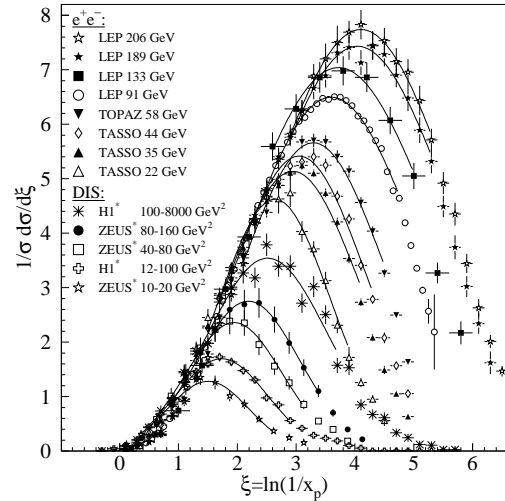
$$xD(x, s) \propto \exp\left[-\frac{1}{2\sigma^2}(\xi - \xi_p)^2\right], \quad (20.9)$$

with the peak position and width varying with the energy as [67] (see also Ref. [2])

$$\xi_p \simeq \frac{1}{4} \ln\left(\frac{s}{\Lambda^2}\right), \quad \sigma \propto \left[\ln\left(\frac{s}{\Lambda^2}\right)\right]^{3/4}. \quad (20.10)$$

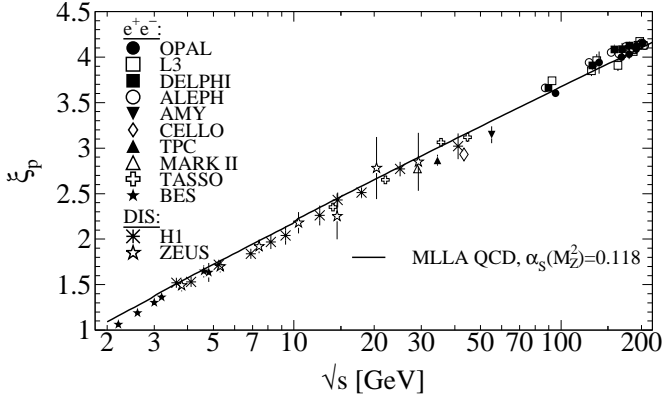
Next-to-leading logarithmic corrections to the above predictions have been calculated [68]. In the method of Ref. [69], see also Refs. [70,71], the corrections are included in an analytical form known as the ‘modified leading logarithmic approximation’ (MLLA). Alternatively they can be used to compute higher-moment corrections to the shape in Eq. (20.9) [72]. The small- $x$  resummation of the coefficient functions for semi-inclusive  $e^+e^-$  annihilation and the timelike splitting functions in the standard  $\overline{\text{MS}}$  scheme was recently extended in Refs. [73,74] and has reached fully analytic next-to-next-to-leading logarithmic accuracy. First applications of these results to gluon and quark jet multiplicities have been presented in Refs. [75].

Fig. 20.4 shows the  $\xi$  distribution for charged particles produced in the current region of the Breit frame in DIS and in  $e^+e^-$  annihilation. Consistent with Eq. (20.9) (the ‘hump backed plateau’) and Eq. (20.10) the distributions have a Gaussian shape with the peak position and area increasing with the CM energy ( $e^+e^-$ ) and  $Q^2$  (DIS).



**Figure 20.4:** Distribution of  $\xi = \ln(1/x_p)$  at several CM energies ( $e^+e^-$ ) [28–29,34–37,76–79] and intervals of  $Q^2$  (DIS) [58,59]. At each energy only one representative measurement is displayed. For clarity some measurements at intermediate CM energies ( $e^+e^-$ ) or  $Q^2$  ranges (DIS) are not shown. The DIS measurements (\*) have been scaled by a factor of 2 for direct comparability with the  $e^+e^-$  results. Fits of simple Gaussian functions are overlaid for illustration.

The predicted energy dependence Eq. (20.10) of the peak in the  $\xi$  distribution is explained by soft gluon coherence (angular ordering), *i.e.*, the destructive interference of the color wavefunction of low energy gluon radiation, which correctly predicts the suppression of hadron production at small  $x$ . Of course, a decrease at very small  $x$  is expected on purely kinematical grounds, but this would occur at particle energies proportional to their masses, *i.e.*, at  $x \propto m/\sqrt{s}$  and hence  $\xi \sim \frac{1}{2} \ln s$ . Thus, if the suppression were purely kinematic, the peak position  $\xi_p$  would vary twice as rapidly with the energy, which is ruled out by the data in Fig. 20.5. The  $e^+e^-$  and DIS data agree well with each other, demonstrating the universality of hadronization, and the MLLA prediction. Measurements of the higher moments of the  $\xi$  distribution in  $e^+e^-$  [37,79–81] and DIS [59] have also been performed and show consistency with each other.



**Figure 20.5:** Evolution of the peak position,  $\xi_p$ , of the  $\xi$  distribution with the CM energy  $\sqrt{s}$ . The MLLA QCD prediction using  $\alpha_S(s = M_Z^2) = 0.118$  is superimposed to the data of Refs. [28–30,33–37,57,58,77,78,81–89].

The average charged particle multiplicity is another observable sensitive to fragmentation functions for small particle momenta. Perturbative predictions using both NLO [90] and MLLA [91,93] have been obtained from solving Eq. (20.4) yielding

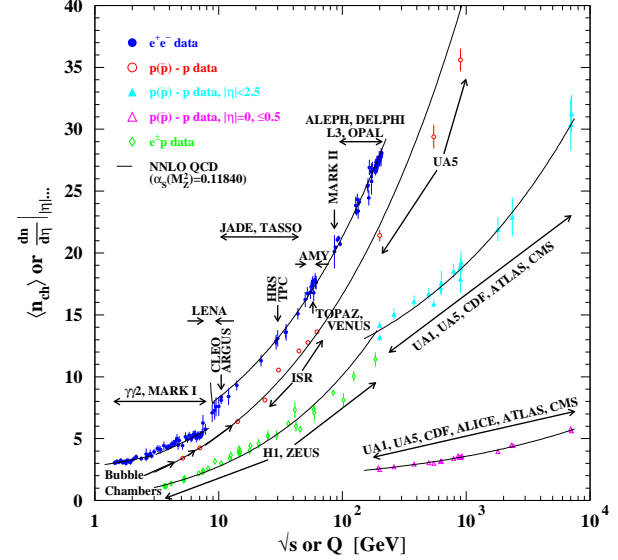
$$\langle n_G(Q^2) \rangle \propto \alpha_S^b(Q^2) \cdot \exp \left[ \frac{c}{4\pi b_0 \sqrt{\alpha_S(Q^2)}} \cdot \left( 1 + 6a_2 \frac{\alpha_S(Q^2)}{\pi} \right) \right] \quad (20.11)$$

where  $b = \frac{1}{4} + \frac{10}{27} \frac{n_f}{4\pi b_0}$ ,  $c = \sqrt{96\pi}$ , with  $b_0 = (33 - 2n_f)/(12\pi)$ , cp. Section 9 of this *Review*, for  $n_f$  contributing quark flavours. Higher order corrections to Eq. (20.11) are known up to next-to-next-to-next-to-leading order (3NLO), for details and references see [94]. The term proportional to  $a_2 \approx -0.502 + 0.0421 n_f - 0.00036 n_f^2$  in Eq. (20.11) is the contribution due to NNLO corrections [95]. The quantity  $\langle n_G(Q^2) \rangle$  strictly refers to the average number of gluons, while for quarks a correction factor  $r = \langle n_G \rangle / \langle n_q \rangle$  weakly depending on  $Q^2$  is required due to the different color factors in quark and gluon couplings, respectively. Higher order corrections up to 3NLO on the asymptotic value  $r = C_A/C_F = 9/4$  [96] are quoted in [94].

Employing the hypothesis of ‘Local Parton-Hadron Duality’ (LPHD) [91], *i.e.*, that the color charge of partons is balanced locally in phase space and, hence, their hadronization occurs locally such that (Mellin transformed) parton and hadron inclusive distributions directly correspond, Eq. (20.11) can be applied to describe average charged particle multiplicities obtained in  $e^+e^-$  annihilation. The equation can also be applied to  $e^\pm p$  scattering if the current fragmentation region of the Breit frame is considered for measuring the average charged particle multiplicity. Fig. 20.6 shows corresponding data and fits of Eq. (20.11) where apart from a LPHD normalization factor a constant offset has been allowed for, that is  $\langle n_{ch}(Q) \rangle = K_{LHPD} \cdot \langle n_G(Q) \rangle / r + n_0$ .

In hadron-hadron collisions beam remnants, *e.g.* from single-diffractive (SD) scattering where one colliding proton is negligibly deflected while hadrons are related with the other colliding proton are well-separated in rapidity from the former proton, contribute to the measurement of the hadron multiplicity from a hard parton-parton scattering, making interpretation of the data more model dependent. Experimental results are usually given for inelastic processes or for non-single diffractive processes (NSD). Due to the large beam particle momenta at Tevatron and LHC, not all final state particles can be detected within the limited detector acceptance. Therefore, experiments at Tevatron and LHC quote particle multiplicities for limited ranges of pseudo-rapidity  $\eta = -\ln \tan(\vartheta/2)$  or at central rapidity, *i.e.*  $\eta = 0$ , shown in Fig. 20.6.

An universality of the average particle multiplicities in  $e^+e^-$  and  $p(\bar{p})$  processes has been reported in Ref. [122] when considering an effective collision energy  $Q_{\text{eff}} = \sqrt{s}/k$  in  $p(\bar{p})$  reduced by a factor of  $k \approx 3$  plus a constant offset of  $n_0 \approx 2$ . A more detailed review is available in Ref. [123]. According to investigations presented



**Figure 20.6:** Average charged particle multiplicity  $\langle n_{ch} \rangle$  as a function of  $\sqrt{s}$  or  $Q$  for  $e^+e^-$  and  $p\bar{p}$  annihilations, and  $pp$  and  $ep$  collisions. The indicated errors are statistical and systematic uncertainties added in quadrature, except when no systematic uncertainties are given. All NNLO QCD curves are Eq. (20.11) with fitted normalization,  $K_{LHPD}$ , and offset,  $n_0$ , using a fixed  $\alpha_S(M_Z^2) = 0.1184$  [92] and for  $e^+e^-$  annihilation data  $n_f = 3, 4$ , or  $5$  depending on  $\sqrt{s}$ , else  $n_f = 3$ .  $e^+e^-$ : Contributions from  $K_S^0$  and  $\Lambda$  decays included. Data compiled from Refs. [8,9,28,34,35,40,78,84,97–107].  $e^\pm p$ : Multiplicities have been measured in the current fragmentation region of the Breit frame. Data compiled from Refs. [58,59,63,108,109].  $p(\bar{p})$ : Measured values above 20 GeV refer to non-single diffractive (NSD) processes. Central pseudorapidity multiplicities  $(dn/d\eta)|_{\eta=0}$  refer to either  $|\eta| < 2.5$  (CMS:  $|\eta| < 2.4$ ) or  $|\eta| = 0$  (UA5, CMS, ALICE:  $|\eta| < 0.5$ ). Data compiled from Refs. [110–121].

in Ref. [124] the universality of the energy dependence of average particle multiplicities also applies to hadron-hadron and nucleus-nucleus collisions for both full and central rapidity multiplicities. Evidence for this universality is given by the good agreement for the energy dependence of Eq. (20.11) when fit to the  $p(\bar{p})$  data as shown in Fig. 20.6.

## 20.4. Fragmentation models

Although the scaling violation can be calculated perturbatively, the actual form of the parton fragmentation functions is non-perturbative. Perturbative evolution gives rise to a shower of quarks and gluons (partons). Multi-parton final states from leading and higher order matrix element calculations are linked to these parton showers using factorization prescriptions, also called matching schemes, see Ref. [125] for an overview. Phenomenological schemes are then used to model the carry-over of parton momenta and flavor to the hadrons. Implemented in Monte Carlo event generators (see Section 41 of this *Review*), these schemes have been tuned using  $e^+e^-$  data and provide good description of hadron collisions as well, thus providing evidence of the universality of the fragmentation functions.

## 20.5. Quark and gluon fragmentation functions

The fragmentation functions are solutions to the evolution equations Eq. (20.4), but need to be parametrized at some initial scale  $\mu_0^2$  (usually around 1 GeV<sup>2</sup> for light quarks and gluons and  $m_Q^2$  for heavy quarks). A usual parametrization for light hadrons is [134–141]

$$D_i^h(x, \mu_0^2) = N x^\alpha (1-x)^\beta \left( 1 + \gamma(1-x)^\delta \right), \quad (20.12)$$

where the normalization  $N$ , and the parameters  $\alpha$ ,  $\beta$ ,  $\gamma$  and  $\delta$  in general depend on the energy scale  $\mu_0^2$ , and also on the type of the parton,  $i$ , and the hadron,  $h$ . Frequently the term involving  $\gamma$  and  $\delta$  is left out [136–139]. Heavy flavor fragmentation into heavy mesons is discussed in Sec. 20.9. The parameters of Eq. (20.12) (see [134–139]) are obtained by performing global fits to data on various hadron types for different combinations of partons and hadrons in  $e^+e^-$ , lepton-hadron and hadron-hadron collisions.

Sets of fragmentation functions are available for pions, kaons, protons, neutrons, etas, Lambdas and charged hadrons [134–141].

Data from  $e^+e^-$  annihilation present the cleanest experimental source for the measurement of fragmentation functions, but can not contribute to disentangle quark from antiquark distributions. Since the bulk of the  $e^+e^-$  annihilation data is obtained at the mass of the  $Z$ -boson, where the electroweak couplings are roughly the same for the different partons, it provides the most precise determination of the flavor-singlet quark fragmentation. Flavor tagged results [142], distinguishing between the light quark, charm and bottom contributions are of particular value for flavor decomposition, even though those measurements can not be unambiguously interpreted in perturbative QCD.

The most relevant source for quark-antiquark (and also flavor) separation is provided by data from semi-inclusive DIS (SIDIS). Semi-inclusive measurements are usually performed at much lower scales than for  $e^+e^-$  annihilation. The inclusion of SIDIS data in global fits allows for a wider coverage in the evolution of the fragmentation functions, resulting at the same time in a stringent test of the universality of these distributions. Charged-hadron production data in hadronic collisions also presents a sensitivity on (anti-)quark fragmentation functions.

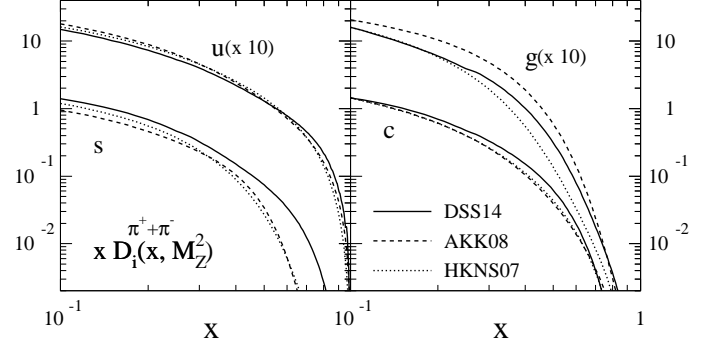
The gluon fragmentation function  $D_g(x)$  can be extracted, in principle, from the longitudinal fragmentation function  $F_L$  in Eq. (20.2), as the coefficient functions  $C_{L,i}$  for quarks and gluons are comparable at order  $\alpha_s$ . However at NLO, *i.e.*, including the  $\mathcal{O}(\alpha_s^2)$  coefficient functions  $C_{L,i}^{(2)}$  [25], quark fragmentation is dominant in  $F_L$  over a large part of the kinematic range, reducing the sensitivity on  $D_g$ . This distribution could be determined also analyzing the evolution of the fragmentation functions. This possibility is limited by the lack of sufficiently precise data at energy scales away from the  $Z$ -resonance and the dominance of the quark contributions and at medium and large values of  $x$ .

$D_g$  can also be deduced from the fragmentation of three-jet events in which the gluon jet is identified, for example, by tagging the other two jets with heavy quark decays. To leading order, the measured distributions of  $x = E_{\text{had}}/E_{\text{jet}}$  for particles in gluon jets can be identified directly with the gluon fragmentation function  $D_g(x)$ . At higher orders the theoretical interpretation of this observable is ambiguous.

A comparison of recent fits of NLO fragmentation functions for  $\pi^+ + \pi^-$  obtained by DSS14 [141], AKK08 [135] and HKNS07 [139] is shown in Fig. 20.7. Differences between the sets are large especially for the gluon fragmentation function over the full range of  $x$  and for the quark distribution at large momentum fractions. The differences are even larger for other species of hadrons like kaons and protons [134,135,139]. Recent analyses [139,141,143] estimate the uncertainties involved in the extraction of fragmentation functions.

A direct constraint on  $D_g$  is provided by  $pp, p\bar{p} \rightarrow hX$  data. At variance with  $e^+e^-$  annihilation and SIDIS, for this process gluon fragmentation starts to contribute at the lowest order in the coupling constant, introducing a strong sensitivity on  $D_g$ . At large  $x \gtrsim 0.5$ , where information from  $e^+e^-$  is sparse, data from hadronic colliders facilitate significantly improved extractions of  $D_g$  [134,135,141]. Recent LHC data has been included in the latest update for pion-fragmentation functions in [141], see Sec.(17.7) for more details.

Photonic fragmentation functions play a relevant role in the theoretical understanding of inclusive photon production in (leptonic and hadronic) high energy processes. Similar to the analogy of parton fragmentation functions and parton distributions in deep inelastic



**Figure 20.7:** Comparison of up, strange, charm and gluon NLO fragmentation functions for  $\pi^+ + \pi^-$  at the mass of the  $Z$ . The different lines correspond to the result of the most recent analyses performed in Refs. [135,139,141].

scattering, also photonic fragmentation functions are analogous to the photon structure function  $F_2^\gamma$  (see review on structure functions in Section 19 of this Review). Since photons have a pointlike coupling to quarks [144], the corresponding fragmentation functions obey inhomogeneous evolution equations and are generally decomposed into a perturbative and a non-perturbative component [138,145,146]. The hadronic part, sometimes approximated by the Vector Meson Dominance Model, can be obtained by performing global analysis to the available prompt photon data [7,30,33,37–39,86,147,179].

## 20.6. Identified particles in $e^+e^-$ and semi-inclusive DIS

A great wealth of measurements of  $e^+e^-$  fragmentation into identified particles exists. A collection of references for data on fragmentation into identified particles is given on Table 51.1 of this Review. Representative of this body of data is Fig. 20.8 which shows fragmentation functions as the scaled momentum spectra of charged particles at several CM energies.

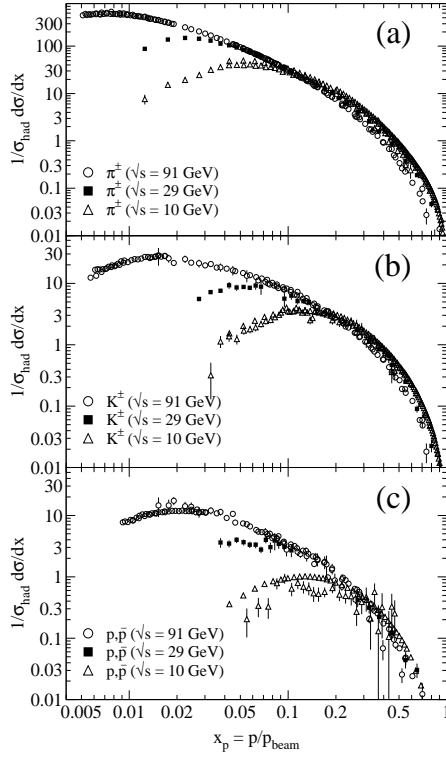
Quantitative results of studies of scaling violation in  $e^+e^-$  fragmentation have been reported in [7,39,149,150]. The values of  $\alpha_s$  obtained are consistent with the world average (see review on QCD in Section 9 of this Review).

Many studies have been made of identified particles produced in lepton-hadron scattering, although fewer particle species have been measured than in  $e^+e^-$  collisions. References [151–158] and [159–165] are representative of the data from fixed target and  $ep$  collider experiments, respectively.

QCD calculations performed at NLO provide an overall good description of the HERA data [54,55,59,165–167] for both SIDIS [168] and the hadron transverse momentum distribution [50] in the kinematic regions in which the calculations are predictive.

Fig. 20.9(a) compares lower-energy fixed-target and HERA data on strangeness production, showing that the HERA spectra have substantially increased multiplicities, albeit with insufficient statistical precision to study scaling violations. The fixed-target data show that the  $\Lambda$  rate substantially exceeds the  $\bar{\Lambda}$  rate in the remnant region, owing to the conserved baryon number from the baryon target. Fig. 20.9(b) shows neutral and charged pion fragmentation functions  $1/N \cdot dn/dz$ , where  $z$  is defined as the ratio of the pion energy to that of the exchanged boson, both measured in the laboratory frame. Results are shown from HERMES and the EMC experiments, where HERMES data have been evolved with NLO QCD to  $\langle Q^2 \rangle = 25 \text{ GeV}^2$  in order to be consistent with the EMC. Each of the experiments uses various kinematic cuts to ensure that the measured particles lie in the region which is expected to be associated with the struck quark. In the DIS kinematic regime accessed at these experiments, and over the range in  $z$  shown in Fig. 20.9, the  $z$  and  $x_F$  variables have similar values [51]. The precision data on identified particles can be used in the study of the quark flavor content of the proton [169].

Data on identified particle production can aid the investigation of the universality of jet fragmentation in  $e^+e^-$  and DIS. The strangeness



**Figure 20.8:** Scaled momentum spectra of (a)  $\pi^\pm$ , (b)  $K^\pm$ , and (c)  $p/\bar{p}$  at  $\sqrt{s} = 10, 29$ , and  $91$  GeV [42–44,86,147,148].

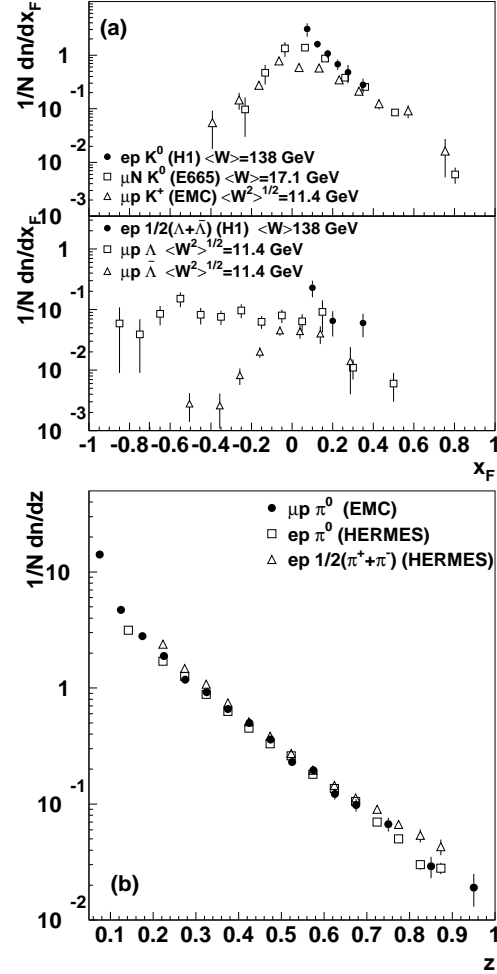
suppression factor  $\gamma_s$ , as derived principally from tuning the Lund string model [127] within JETSET [128], is typically found to be around 0.3 in  $e^+e^-$  experiments [76], although values closer to 0.2 [170] have also been obtained. A number of measurements of so-called  $V^0$ -particles ( $K^0$ ,  $\Lambda^0$ ) and the relative rates of  $V^0$ 's and inclusively produced charged particles have been performed at HERA [159–161] and fixed target experiments [151]. These typically favour a stronger suppression ( $\gamma_s \approx 0.2$ ) than usually obtained from  $e^+e^-$  data although values close to 0.3 have also been obtained [171,172].

However, when comparing the description of QCD-based models for lepton-hadron interactions and  $e^+e^-$  collisions, it is important to note that the overall description by event generators of inclusively produced hadronic final states is more accurate in  $e^+e^-$  collisions than lepton-hadron scattering are affected by uncertainties in the modelling of the parton composition of the proton and photon, the extended target remnant, and initial and final-state QCD radiation. Furthermore, the tuning of event generators for  $e^+e^-$  collisions is typically based on a larger set of parameters and uses more observables [76] than are used when optimizing models for lepton-hadron data [174].

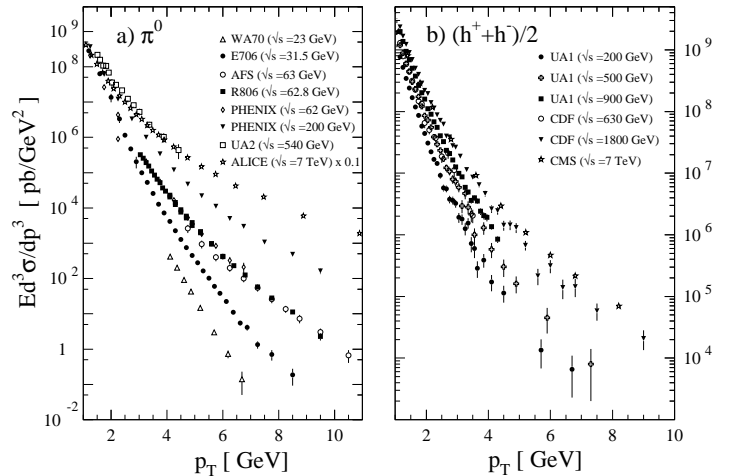
## 20.7. Fragmentation in hadron-hadron collisions

An extensive set on high-transverse momentum ( $p_T$ ) single-inclusive hadron data has been collected in  $h_1 h_2 \rightarrow hX$  scattering processes, both at high energy colliders and fixed-target experiments [175–194]. Only the transverse momentum  $p_T$  is considered in hadron-hadron collisions because of lack of knowledge of the longitudinal momentum of the hard subprocess. Fig. 20.10 shows the cross section (which is proportional to the particle number) density  $\frac{d^3\sigma}{dp^3} = \frac{d^3\sigma}{dp_x dp_y dp_z} = \frac{E}{\pi m^2} \frac{d^2\sigma}{dy d(p_T^2)}$  for a compilation of neutral pion and charged hadron production data for energies in the range  $\sqrt{s} \approx 23$  – 7000 GeV. More data for different hadron species has been recently obtained at high energy colliders [195–199].

The differential cross-section for high-transverse momentum distributions has been computed to next-to-leading order accuracy in



**Figure 20.9:** (a)  $1/N \cdot dn/dx_F$  for identified strange particles in DIS at various values of  $W$  [151,154,159]. (b)  $1/N \cdot dn/dz$  for measurements of pions from fixed-target DIS experiment [152,155,158].



**Figure 20.10:** Selection of inclusive (a)  $\pi^0$  and (b) charged-hadron production data from  $pp$  [120,183,187,191–194] and  $p\bar{p}$  [175,178,181] collisions.

perturbative QCD [200]. The factorization,  $\mu_f$ , and renormalization,  $\mu$ , scales of these calculations typical range from  $p_T^2/4 \leq \mu_f^2, \mu^2 \leq 4p_T^2$ . NLO calculations significantly under-predict the cross-section for

several fixed-target energy data sets [201,202]. Different strategies have been developed to ameliorate the theoretical description at fixed-target energies. A possible phenomenological approach involves the introduction of a non-perturbative intrinsic partonic transverse momentum [194,203,204]. From the perturbative side, the resummation of the dominant higher order corrections at threshold produces an enhancement of the theoretical calculation that significantly improves the description of the data [205,206].

Data collected at high energy colliders are either included in global fit analyses or used as a test for the universality of fragmentation functions. Certain tension has been observed between data sets from lower-energy (RHIC) and higher-energy (LHC) collisions [207]. The tension can be largely resolved by excluding from the analysis data with transverse momentum smaller than  $\approx 5 - 10$  GeV, where fixed order pQCD calculations are not expected to provide an accurate description of the process. Still, after removing the smallish  $p_T$  values where the data sets appear to be mutually exclusive in the global fit, lower-energy collisions data show a preference towards harder gluon fragmentation at large  $z$  than LHC data [141].

Measurements of hadron production in longitudinally polarized  $pp$  collisions are used mainly in the determination of the polarized gluon distribution in the proton [208,209].

Hadron production provides a critical observable for probing the high energy-density matter produced in heavy-ion collisions. Measurements at colliders show a suppression of inclusive hadron yields at high transverse momentum for  $AA$  collisions compared to  $pp$  scattering, indicating the formation of a dense medium opaque to quark and gluons, see e.g. [210].

## 20.8. Spin-dependent fragmentation

Measurements of charged-hadron production in unpolarized lepton-hadron scattering provide a unique tool to perform a flavor-separation determination of polarized parton densities from DIS interactions with longitudinally polarized targets [211–215].

Polarized scattering presents the possibility to measure the spin transfer from the struck quark to the final hadron, and thus develop spin-dependent fragmentation functions [216,217]. Early measurements of the longitudinal spin transfer to Lambda hyperons have been presented in [218,219]. This process is also useful in the study of the quark transversity distribution [220], which describes the probability of finding a transversely polarized quark with its spin aligned or anti-aligned with the spin of a transversely polarized nucleon. The transversity function is chiral-odd, and therefore not accessible through measurements of inclusive lepton-hadron scattering. Semi-inclusive DIS, in which another chiral-odd observable may be involved, provides a valuable tool to probe transversity. The Collins fragmentation function [221] relates the transverse polarization of the quark to that of the final hadron. It is chiral-odd and naive T-odd, leading to a characteristic single spin asymmetry in the azimuthal angular distribution of the produced hadron in the hadron scattering plane. Azimuthal angular distributions in semi-inclusive DIS can also be produced by other processes requiring non-polarized fragmentation functions, like the Sivers mechanism [222].

A number of experiments have measured these asymmetries [223–233]. Collins and Sivers asymmetries have been shown experimentally to be non zero by the HERMES measurements on transversely polarized proton targets [224–226]. Independent information on the Collins function has been provided by the BELLE Collaboration [227–228]. Measurements performed by the COMPASS collaboration on deuteron targets show results compatible with zero for both asymmetries [229–231].

## 20.9. Heavy quark fragmentation

It was recognized very early [234] that a heavy flavored meson should retain a large fraction of the momentum of the primordial heavy quark, and therefore its fragmentation function should be much harder than that of a light hadron. In the limit of a very heavy quark, one expects the fragmentation function for a heavy quark to go into any heavy hadron to be peaked near  $x = 1$ .

When the heavy quark is produced at a momentum much larger than its mass, one expects important perturbative effects, enhanced by powers of the logarithm of the transverse momentum over the heavy quark mass, to intervene and modify the shape of the fragmentation function. In leading logarithmic order (*i.e.*, including all powers of  $\alpha_s \log m_Q/p_T$ ), the total (*i.e.*, summed over all hadron types) perturbative fragmentation function is simply obtained by solving the leading evolution equation for fragmentation functions, Eq. (20.4), with the initial condition due to the finite mass of the heavy quark given by  $D_Q(z, \mu^2)|_{\mu^2=m_Q^2} = \delta(1-z)$  and  $D_i(z, \mu^2)|_{\mu^2=m_Q^2} = 0$  for  $i \neq Q$  (here  $D_i(z, \mu^2)$ , stands for the probability to produce a heavy quark  $Q$  from parton  $i$  with a fraction  $z$  of the parton momentum).

Several extensions of the leading logarithmic result have appeared in the literature. Next-to-leading-log (NLL) order results for the perturbative heavy quark fragmentation function have been obtained in [235]. The resummation of the dominant logarithmic contributions at large  $z$  was performed in [45] to next-to-leading-log accuracy. Fixed-order calculations of the fragmentation function at order  $\alpha_s^2$  in  $e^+e^-$  annihilation have appeared in [236] while the initial condition for the perturbative heavy quark fragmentation function has been extended to NNLO in [237].

Inclusion of non-perturbative effects in the calculation of the heavy-quark fragmentation function is done by convoluting the perturbative result with a phenomenological non-perturbative form. This form follows from the simple kinematical consideration that the formation of a hadron by attaching light quarks/anti-quarks to the heavy quark will slightly decelerate the heavy quark. Thus its shape will show a peak which becomes increasingly centered next to  $z = 1$  the higher the quark mass. Among the most popular parametrizations we have the following:

$$\text{Peterson } et al. [238]: D_{np}(z) \propto \frac{1}{z} \left(1 - \frac{1}{z} - \frac{\epsilon}{1-z}\right)^{-2} \quad (20.13)$$

$$\text{Kartvelishvili } et al. [239]: D_{np}(z) \propto z^\alpha (1-z), \quad (20.14)$$

$$\text{Collins\&Spiller [240]: } D_{np}(z) \propto \left(\frac{1-z}{z} + \frac{(2-z)\epsilon_C}{1-z}\right) \times (1+z^2) \left(1 - \frac{1}{z} - \frac{\epsilon_C}{1-z}\right)^{-2} \quad (20.15)$$

$$\text{Colangelo\&Nason [241]: } D_{np}(z) \propto (1-z)^\alpha z^\beta \quad (20.16)$$

$$\text{Bowler [242]: } D_{np}(z) \propto z^{-(1+bm_{h,\perp}^2)} (1-z)^a \exp\left(-\frac{bm_{h,\perp}^2}{z}\right) \quad (20.17)$$

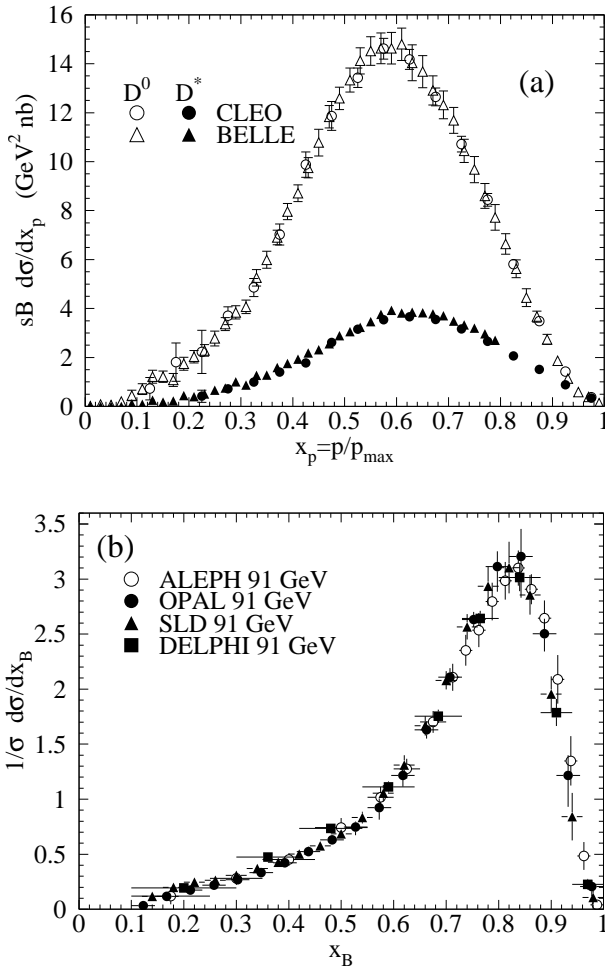
$$\text{Braaten } et al. [243]: \quad (\text{see Eq. (31), (32) in [243]}) \quad (20.18)$$

where  $\epsilon$ ,  $\epsilon_C$ ,  $a$ ,  $bm_{h,\perp}^2$ ,  $\alpha$ , and  $\beta$  are non-perturbative parameters, depending upon the heavy hadron considered. The parameters entering the non-perturbative forms are fitted together with some model of hard radiation, which can be either a shower Monte Carlo, a leading-log or NLL calculation (which may or may not include Sudakov resummation), or a fixed order calculation. In [236], for example, the Peterson *et al.* [238]  $\epsilon$  parameter for charm and bottom production is fitted from the measured distributions of refs. [244,257] for charm, and of [262] for bottom. If the leading-logarithmic approximation (LLA) is used for the perturbative part, one finds  $\epsilon_c \approx 0.05$  and  $\epsilon_b \approx 0.006$ ; if a second order calculation is used one finds  $\epsilon_c \approx 0.035$  and  $\epsilon_b \approx 0.0033$ ; if a NLL improved fixed order  $\mathcal{O}(\alpha_s^2)$  calculation is used instead of NLO  $\mathcal{O}(\alpha_s)$  one finds  $\epsilon_c \approx 0.022$  and  $\epsilon_b \approx 0.0023$ . The larger values found in the LL approximation are consistent with what is obtained in the context of parton shower models [246], as expected. The  $\epsilon$  parameter for charm and bottom scales roughly with the inverse square of the heavy flavour mass. This behaviour can be justified by several arguments [234,247,248]. It can be used to relate the non-perturbative parts of the fragmentation functions of charm and bottom quarks [236,241,249].

A more conventional approach [250] involves the introduction of a unique set of heavy quark fragmentation functions of non-perturbative nature that obey the usual massless evolution equations in Eq. (20.4).

Finite mass terms of the form  $(m_Q/p_T)^n$  are kept in the corresponding short distance coefficient function for each scattering process. Within this approach, the initial condition for the perturbative fragmentation function provides the term needed to define the correct subtraction scheme to match the massless limit for the coefficient function (see e.g. [251]). Such implementation is in line with the variable flavor number scheme introduced for parton distributions functions, as described in Section 19 of this *Review*.

High statistics data for charmed mesons production near the  $\Upsilon$  resonance (excluding decay products of  $B$  mesons) have been published [252,253]. They include results for  $D$  and  $D^*$ ,  $D_s$  (see also [254,255]) and  $\Lambda_c$ . Shown in Fig. 20.11(a) are the CLEO and BELLE inclusive cross-sections times branching ratio  $\mathcal{B}$ ,  $s \cdot \mathcal{B} d\sigma/dx_p$ , for the production of  $D^0$  and  $D^{*+}$ . The variable  $x_p$  approximates the light-cone momentum fraction  $z$ , but is not identical to it. The two measurements are consistent with each other.



**Figure 20.11:** (a) Efficiency-corrected inclusive cross-section measurements for the production of  $D^0$  and  $D^{*+}$  in  $e^+e^-$  measurements at  $\sqrt{s} \approx 10.6$  GeV, excluding  $B$  decay products [252,253]. (b) Measured  $e^+e^-$  fragmentation function of  $b$  quarks into  $B$  hadrons at  $\sqrt{s} \approx 91$  GeV [263].

The branching ratio  $\mathcal{B}$  represents  $D^0 \rightarrow K^-\pi^+$  for the  $D^0$  results and for the  $D^{*+}$  the product branching fraction:  $D^{*+} \rightarrow D^0\pi^+$ ,  $D^0 \rightarrow K^-\pi^+$ . Given the high precision of CLEO's and BELLE's data, a superposition of different parametric forms for the non-perturbative contribution is needed to obtain a good fit [23]. Older studies are reported in Refs. [256–258]. Charmed meson spectra on the  $Z$  peak have been published by OPAL and ALEPH [133,259].

Charm quark production has also been extensively studied at HERA by the H1 and ZEUS collaborations. Measurements have been

made of  $D^{*\pm}$ ,  $D^\pm$ , and  $D_s^\pm$  mesons and the  $\Lambda_c$  baryon. See, for example, Refs. [260,261].

Experimental studies of the fragmentation function for  $b$  quarks, shown in Fig. 20.11(b), have been performed at LEP and SLD [262–264]. Commonly used methods identify the  $B$  meson through its semileptonic decay or based upon tracks emerging from the  $B$  secondary vertex. Heavy flavour contributions from gluon splitting are usually explicitly removed before fitting for the fragmentation functions. The studies in [263] fit the  $B$  spectrum using a Monte Carlo shower model supplemented with non-perturbative fragmentation functions yielding consistent results.

The experiments measure primarily the spectrum of  $B$  mesons. This defines a fragmentation function which includes the effect of the decay of higher mass excitations, like the  $B^*$  and  $B^{**}$ . In the literature (cf. details in Ref. [266]), there is sometimes ambiguity in what is defined to be the bottom fragmentation function. Instead of using what is directly measured (i.e., the  $B$  meson spectrum) corrections are applied to account for  $B^*$  or  $B^{**}$  production in some cases.

Heavy-flavor production in  $e^+e^-$  collisions is the primary source of information for the role of fragmentation effects in heavy-flavor production in hadron-hadron and lepton-hadron collisions. The QCD calculations tend to underestimate the data in certain regions of phase space. Some experimental results from LHC summarized in [267] show such deviations e.g. at high transverse jet momentum and also at low di-jet separation angles, see [268] for details, and were already theoretically investigated in [269].

Both bottomed- and charmed-mesons spectra have been measured at the Tevatron with unprecedented accuracy [270]. The measured spectra are in good agreement with QCD calculations (including non-perturbative fragmentation effects inferred from  $e^+e^-$  data [271]).

The HERA collaborations have produced a number of measurements of beauty production; see, for example, Refs. [260,272–275]. As for the Tevatron data, the HERA results are described well by QCD-based calculations using fragmentation models optimised with  $e^+e^-$  data.

Besides degrading the fragmentation function by gluon radiation, QCD evolution can also generate soft heavy quarks, increasing in the small  $x$  region as  $\sqrt{s}$  increases. Several theoretical studies are available on the issue of how often  $b\bar{b}$  or  $c\bar{c}$  pairs are produced indirectly, via a gluon splitting mechanism [276–278]. Experimental results from studies on charm and bottom production via gluon splitting, given in [259,279–283], yield weighted averages of  $\pi_{g \rightarrow c\bar{c}} = 3.05 \pm 0.45\%$  and  $\pi_{g \rightarrow b\bar{b}} = 0.277 \pm 0.072\%$ , respectively.

#### References:

1. G. Altarelli, Phys. Reports **81**, 1 (1982).
2. R.K. Ellis *et al.*, *QCD and Collider Physics*, Cambridge University Press (1996).
3. S. Albino *et al.*, arXiv:0804.2021 (2008).
4. F. Arleo, Eur. Phys. J. **C61**, 603 (2009).
5. P. Nason and B.R. Webber, Nucl. Phys. **B421**, 473 (1994); Erratum *ibid.* **B480**, 755 (1996).
6. J.D. Bjorken and E.A. Paschos, Phys. Rev. **185**, 1975 (1969); R.P. Feynman, *Photon Hadron Interactions*, Benjamin, New York (1972).
7. ALEPH Collab.: D. Barate *et al.*, Phys. Lett. **B357**, 487 (1995); Erratum *ibid.*, **B364**, 247 (1995).
8. OPAL Collab.: R. Akers *et al.*, Z. Phys. **C68**, 203 (1995).
9. DELPHI Collab.: P. Abreu *et al.*, Eur. Phys. J. **C6**, 19 (1999).
10. W. Kittel and E.A. De Wolf, *Soft Multihadron Dynamics*, World Scientific (2005).
11. H.F. Jones, Nuovo Cimento **40A**, 1018 (1965); K.H. Streng *et al.*, Z. Phys. **C2**, 237 (1979).
12. V.N. Gribov and L.N. Lipatov, Sov. J. Nucl. Phys. **15**, 438 (1972); V.N. Gribov and L.N. Lipatov, Sov. J. Nucl. Phys. **15**, 675 (1972); L.N. Lipatov, Sov. J. Nucl. Phys. **20**, 95 (1975); G. Altarelli and G. Parisi, Nucl. Phys. **B126**, 298 (1977); Yu.L. Dokshitzer, Sov. Phys. JETP Lett. **46**, 641 (1977).



13. H. Georgi and H.D. Politzer, Nucl. Phys. **B136**, 445 (1978);  
J.F. Owens, Phys. Lett. **B76**, 85 (1978);  
T. Uematsu, Phys. Lett. **B79**, 97 (1978).
14. G. Curci *et al.*, Nucl. Phys. **B175**, 27 (1980).
15. W. Furmanski and R. Petronzio, Phys. Lett. **97B**, 437 (1980).
16. E.G. Floratos *et al.*, Nucl. Phys. **B192**, 417 (1981);  
T. Muehisa *et al.*, Prog. Theor. Phys. **67**, 609 (1982).
17. J. Kalinowski *et al.*, Nucl. Phys. **B181**, 221 (1981);  
J. Kalinowski *et al.*, Nucl. Phys. **B181**, 253 (1981).
18. M. Stratmann and W. Vogelsang, Nucl. Phys. **B496**, 41 (1997).
19. Yu.L. Dokshitzer *et al.*, Phys. Lett. **B634**, 504 (2006).
20. A. Mitov *et al.*, Phys. Lett. **B638**, 61 (2006).
21. S. Moch and A. Vogt, Phys. Lett. **B659**, 290 (2008).
22. A.A. Almasy, A. Vogt, S. Moch, Nucl. Phys. **B854**, 133 (2013).
23. M. Cacciari *et al.*, JHEP **0604**, 006 (2006);  
M. Cacciari *et al.*, JHEP **0510**, 034 (2005).
24. G. Altarelli *et al.*, Nucl. Phys. **B160**, 301 (1979);  
R. Baier and K. Fey, Z. Phys. **C2**, 339 (1979).
25. P.J. Rijken and W.L. van Neerven, Phys. Lett. **B386**, 422 (1996);  
P.J. Rijken and W.L. van Neerven, Phys. Lett. **B392**, 207 (1997);  
P.J. Rijken and W.L. van Neerven, Nucl. Phys. **B487**, 233 (1997).
26. A. Mitov and S. Moch, Nucl. Phys. **B751**, 18 (2006).
27. ALEPH Collab.: E. Barate *et al.*, Phys. Reports **294**, 1 (1998).
28. ALEPH Collab.: D. Buskulic *et al.*, Z. Phys. **C73**, 409 (1997).
29. L3 Collab.: B. Adeva *et al.*, Phys. Lett. **B259**, 199 (1991).
30. AMY Collab.: Y.K. Li *et al.*, Phys. Rev. **D41**, 2675 (1990).
31. HRS Collab.: D. Bender *et al.*, Phys. Rev. **D31**, 1 (1984).
32. MARK II Collab.: G.S. Abrams *et al.*, Phys. Rev. Lett. **64**, 1334 (1990).
33. MARK II Collab.: A. Petersen *et al.*, Phys. Rev. **D37**, 1 (1988).
34. OPAL Collab.: R. Akers *et al.*, Z. Phys. **C72**, 191 (1996).
35. OPAL Collab.: K. Ackerstaff *et al.*, Z. Phys. **C75**, 193 (1997).
36. OPAL Collab.: G. Abbiendi *et al.*, Eur. Phys. J. **C16**, 185 (2000).
37. TASSO Collab.: W. Braunschweig *et al.*, Z. Phys. **C47**, 187 (1990).
38. OPAL Collab.: K. Ackerstaff *et al.*, Eur. Phys. J. **C7**, 369 (1998);  
OPAL Collab.: G. Abbiendi *et al.*, Eur. Phys. J. **C27**, 467 (2003).
39. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B398**, 194 (1997).
40. OPAL Collab.: G. Abbiendi *et al.*, Eur. Phys. J. **C37**, 25 (2004).
41. TASSO Collab.: R. Brandelik *et al.*, Phys. Lett. **B114**, 65 (1982).
42. SLD Collab.: K. Abe *et al.*, Phys. Rev. **D69**, 072003 (2004).
43. TPC Collab.: H. Aihara *et al.*, Phys. Rev. Lett. **61**, 1263 (1988).
44. BELLE Collab.: M. Leitgab *et al.*, Phys. Rev. Lett. **111**, 062002 (2013);  
BaBar Collab.: J.P. Lees *et al.*, Phys. Rev. **D88**, 032011 (2013).
45. M. Cacciari and S. Catani, Nucl. Phys. **B617**, 253 (2001).
46. J. Blümlein and V. Ravindran, Phys. Lett. **B640**, 40 (2006).
47. S. Moch and A. Vogt, Phys. Lett. **B680**, 239 (2009).
48. S. Moch and A. Vogt, JHEP **0911**, 099 (2009).
49. A. Vogt, Phys. Lett. **B691**, 77 (2010).
50. P. Aurenche *et al.*, Eur. Phys. J. **C34**, 277 (2004);  
A. Daleo *et al.*, Phys. Rev. **D71**, 034013 (2005);  
B.A. Kniehl *et al.*, Nucl. Phys. **B711**, 345 (2005);  
Erratum *ibid.* **B720**, 231 (2005).
51. E665 Collab.: M.R. Adams *et al.*, Phys. Lett. **B272**, 163 (1991).
52. EMC Collab.: M. Arneodo *et al.*, Z. Phys. **C35**, 417 (1987).
53. H1 Collab.: I. Abt *et al.*, Z. Phys. **C63**, 377 (1994).
54. ZEUS Collab.: M. Derrick *et al.*, Z. Phys. **C70**, 1 (1996).
55. ZEUS Collab.: J. Breitweg *et al.*, Phys. Lett. **B414**, 428 (1997).
56. H1 Collab.: S. Aid *et al.*, Nucl. Phys. **B445**, 3 (1995).
57. ZEUS Collab.: M. Derrick *et al.*, Z. Phys. **C67**, 93 (1995).
58. H1 Collab.: C. Adloff *et al.*, Nucl. Phys. **B504**, 3 (1997).
59. ZEUS Collab.: J. Breitweg *et al.*, Eur. Phys. J. **C11**, 251 (1999).
60. H1 Collab.: F.D. Aaron *et al.*, Phys. Lett. **B654**, 148 (2007).
61. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B311**, 408 (1993).
62. MARK II Collab.: J.F. Patrick *et al.*, Phys. Rev. Lett. **49**, 1232, (1982).
63. ZEUS Collab.: S. Chekanov *et al.*, JHEP **0806**, 061 (2008).
64. K.H. Streng *et al.*, Z. Phys. **C2**, 237 (1979).
65. A.H. Mueller, Phys. Lett. **B104**, 161 (1981).
66. A. Bassetto *et al.*, Nucl. Phys. **B207**, 189 (1982).
67. Yu.L. Dokshitzer *et al.*, Z. Phys. **C15**, 324 (1982).
68. A.H. Mueller, Nucl. Phys. **B213**, 85 (1983);  
Erratum in *ibid.* **B241**, 141 (1984).
69. Yu.L. Dokshitzer *et al.*, Int. J. Mod. Phys. **A7**, 1875 (1992).
70. Yu.L. Dokshitzer *et al.*, *Basics of Perturbative QCD*, Editions Frontières (1991).
71. V.A. Khoze and W. Ochs, Int. J. Mod. Phys. **A12**, 2949 (1997).
72. C.P. Fong and B.R. Webber, Nucl. Phys. **B355**, 54 (1992).
73. S. Albino *et al.*, Nucl. Phys. **B851**, 86 (2011);  
S. Albino *et al.*, Nucl. Phys. **B855**, 801 (2012).
74. A. Vogt, JHEP **1110**, 025 (2011);  
C.-H. Kom, A. Vogt, K. Yeats, JHEP **1210**, 033 (2012).
75. P. Bolzoni, B.A. Kniehl, A.V. Kotikov, Phys. Rev. Lett. **109**, 242002 (2012);  
P. Bolzoni, B.A. Kniehl, A.V. Kotikov, Nucl. Phys. **B875**, 18 (2013).
76. DELPHI Collab.: P. Abreu *et al.*, Z. Phys. **C73**, 11 (1996).
77. DELPHI Collab.: P. Abreu *et al.*, Z. Phys. **C73**, 229 (1997).
78. L3 Collab.: P. Achard *et al.*, Phys. Reports **399**, 71 (2004).
79. TOPAZ Collab.: R. Itoh *et al.*, Phys. Lett. **B345**, 335 (1995).
80. TASSO Collab.: W. Braunschweig *et al.*, Z. Phys. **C22**, 307 (1990).
81. OPAL Collab.: M.Z. Akrawy *et al.*, Phys. Lett. **B247**, 617 (1990).
82. BES Collab.: J.Z. Bai *et al.*, Phys. Rev. **D69**, 072002 (2004).
83. ALEPH Collab.: D. Buskulic *et al.*, Z. Phys. **C55**, 209 (1992).
84. ALEPH Collab.: A. Heister *et al.*, Eur. Phys. J. **C35**, 457 (2004).
85. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B275**, 231 (1992).
86. DELPHI Collab.: P. Abreu *et al.*, Eur. Phys. J. **C5**, 585 (1998).
87. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B459**, 397 (1999).
88. L3 Collab.: M. Acciarri *et al.*, Phys. Lett. **B444**, 569 (1998).
89. TPC/TWO-GAMMA Collab.: H. Aihara *et al.*, LBL 23737.
90. B.R. Webber, Phys. Lett. **B143**, 501 (1984).
91. Ya.I. Azimov, Yu.L. Dokshitzer, V.A. Khoze, S.I. Troyan, Z. Phys. **C27**, 65 (1985).
92. J. Beringer *et al.* (Particle Data Group), Phys. Rev. **D86**, 010001 (2012).
93. Ya.I. Azimov, Yu.L. Dokshitzer, V.A. Khoze, S.I. Troyan, Z. Phys. **C31**, 213 (1986).
94. I.M. Dremin, J.W. Gary, Phys. Reports **349**, 301 (2001).
95. I.M. Dremin, V.A. Nechitailo, Mod. Phys. Lett. **A9**, 1471 (1994).
96. S.J. Brodsky, J.F. Gunion, Phys. Rev. Lett. **37**, 402 (1976).
97. OPAL Collab.: P.D. Acton *et al.*, Z. Phys. **C53**, 539 (1992) and references therein.
98. ALEPH Collab.: D. Buskulic *et al.*, Z. Phys. **C69**, 15 (1996).
99. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B372**, 172 (1996).
100. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B416**, 233 (1998).
101. DELPHI Collab.: P. Abreu *et al.*, Eur. Phys. J. **C18**, 203 (2000).
102. L3 Collab.: M. Acciarri *et al.*, Phys. Lett. **B371**, 137 (1996).
103. L3 Collab.: M. Acciarri *et al.*, Phys. Lett. **B404**, 390 (1997).
104. L3 Collab.: M. Acciarri *et al.*, Phys. Lett. **B444**, 569 (1998).
105. TOPAZ Collab.: K. Nakabayashi *et al.*, Phys. Lett. **B413**, 447 (1997).
106. VENUS Collab.: K. Okabe *et al.*, Phys. Lett. **B423**, 407 (1998).
107. ARGUS Collab.: H. Albrecht *et al.*, Z. Phys. **C54**, 13 (1992).
108. H1 Collab.: F.D. Aaron *et al.*, Phys. Lett. **B654**, 148 (2007).
109. ZEUS Collab.: S. Chekanov *et al.*, Phys. Lett. **B510**, 36 (2001).
110. J. Benecke *et al.*, Nucl. Phys. **B76**, 29 (1976).

111. W.M. Morse *et al.*, Phys. Rev. **D15**, 66 (1977).
112. W. Thomé *et al.*, Nucl. Phys. **B129**, 365 (1977).
113. A. Breakstone *et al.*, Phys. Rev. **D30**, 528 (1984).
114. UA5 Collab.: G.J. Alner *et al.*, Phys. Reports **154**, 247 (1987).
115. UA5 Collab.: R.E. Ansorge *et al.*, Z. Phys. **C43**, 357 (1989).
116. UA1 Collab.: C. Albajar *et al.*, Nucl. Phys. **B335**, 261 (1990).
117. CDF Collab.: F.Abe *et al.*, Phys. Rev. **D41**, 2330 (1990).
118. ALICE Collab.: K. Aamodt *et al.*, Eur. Phys. J. **C68**, 89 (2010).
119. CMS Collab.: V. Khachatryan *et al.*, JHEP **1002**, 041 (2010).
120. CMS Collab.: V. Khachatryan *et al.*, JHEP **1101**, 079 (2011).
121. CMS Collab.: V. Khachatryan *et al.*, Phys. Rev. Lett. **105**, 022002 (2010).
122. P.V. Chliapnikov, V.A. Uvarov, Phys. Lett. **B251**, 192 (1990).
123. J.F. Grosse-Oetringhaus, K. Reygers, J. Phys. **G37**, 083001 (2010).
124. E.K.G. Sarkisyan, A.S. Sakharov, hep-ph/0410324; E.K.G. Sarkisyan, A.S. Sakharov, AIP Conf. Proc. **828**, 35 (2005), [hep-ph/0510191]; E.K.G. Sarkisyan, A.S. Sakharov, Eur. Phys. J. **C70**, 533 (2010).
125. S. Höche *et al.*, hep-ph/0602031 (2006); J. Alwall *et al.*, Eur. Phys. J. **C53**, 473 (2008); S. Mrenna and P. Richardson, JHEP **0405**, 040 (2004).
126. X. Artru and G. Mennessier, Nucl. Phys. **B70**, 93 (1974).
127. B. Andersson *et al.*, Phys. Reports **97**, 31 (1983).
128. T. Sjöstrand and M. Bengtsson, Comp. Phys. Comm. **43**, 367 (1987); T. Sjöstrand, Comp. Phys. Comm. **82**, 74 (1994).
129. T. Sjöstrand, S. Mrenna, P. Skands, JHEP **0605**, 026 (2006); T. Sjöstrand, S. Mrenna, P. Skands, Comp. Phys. Comm. **178**, 852 (2008).
130. S. Chun and C. Buchanan, Phys. Reports **292**, 239 (1998).
131. G. Marchesini *et al.*, Comp. Phys. Comm. **67**, 465 (1992); G. Corcella *et al.*, JHEP **0101**, 010 (2001); M. Bähr *et al.*, Eur. Phys. J. **C58**, 639 (2008).
132. T. Gleisberg *et al.*, JHEP **0902**, 007 (2009).
133. OPAL Collab.: G. Alexander *et al.*, Z. Phys. **C69**, 543 (1996).
134. D. de Florian *et al.*, Phys. Rev. **D76**, 074033 (2007); D. de Florian *et al.*, Phys. Rev. **D75**, 114010 (2007).
135. S. Albino *et al.*, Nucl. Phys. **B803**, 42 (2008).
136. S. Kretzer *et al.*, Eur. Phys. J. **C22**, 269 (2001).
137. S. Kretzer, Phys. Rev. **D62**, 054001 (2000).
138. L. Bourhis *et al.*, Eur. Phys. J. **C19**, 89 (2001).
139. M. Hirai *et al.*, Phys. Rev. **D75**, 094009 (2007).
140. C. Aidala *et al.*, Phys. Rev. **D83**, 034002 (2011).
141. D. de Florian *et al.*, Phys. Rev. **D91**, 014035 (2015).
142. ALEPH Collab.: R. Barate *et al.*, Eur. Phys. J. **C17**, 1 (2000); OPAL Collab.: R. Akers *et al.*, Z. Phys. **C68**, 179 (1995); OPAL Collab.: G. Abbiendi *et al.*, Eur. Phys. J. **C11**, 217 (1999).
143. M. Epele *et al.*, Phys. Rev. **D86**, 074028 (2012).
144. E. Witten, Nucl. Phys. **210**, 189 (1977).
145. L. Bourhis, M. Fontannaz, and J. P. Guillet, Eur. Phys. J. **C2**, 529 (1998).
146. M. Gluck, E. Reya, and A. Vogt, Phys. Rev. **D48**, 116 (1993); Erratum *ibid.* **D51**, 1427 (1995).
147. SLD Collab.: K. Abe *et al.*, Phys. Rev. **D59**, 052001 (1999).
148. ALEPH Collab.: D. Buskulic *et al.*, Z. Phys. **C66**, 355 (1995); ARGUS Collab.: H. Albrecht *et al.*, Z. Phys. **C44**, 547 (1989); OPAL Collab.: R. Akers *et al.*, Z. Phys. **C63**, 181 (1994).
149. DELPHI Collab.: P. Abreu *et al.*, Eur. Phys. J. **C13**, 573 (2000).
150. B.A. Kniehl *et al.*, Phys. Rev. Lett. **85**, 5288 (2000).
151. E665 Collab.: M.R. Adams *et al.*, Z. Phys. **C61**, 539 (1994).
152. EMC Collab.: J.J. Aubert *et al.*, Z. Phys. **C18**, 189 (1983); EMC Collab.: M. Arneodo *et al.*, Phys. Lett. **B150**, 458 (1985).
153. EMC Collab.: M. Arneodo *et al.*, Z. Phys. **C33**, 167 (1986).
154. EMC Collab.: M. Arneodo *et al.*, Z. Phys. **C34**, 283 (1987).
155. HERMES Collab.: A. Airapetian *et al.*, Eur. Phys. J. **C21**, 599 (2001).
156. HERMES Collab.: A. Airapetian *et al.*, Phys. Rev. **D87**, 074029 (2013).
157. COMPASS Collab.: N. Makke, PoS **DIS2013**, 202 (2013) [arXiv:1307.3407].
158. T.P. McPharlin *et al.*, Phys. Lett. **B90**, 479 (1980).
159. H1 Collab.: S. Aid *et al.*, Nucl. Phys. **B480**, 3 (1996).
160. H1 Collab.: F. D. Aaron *et al.*, Phys. Lett. **B673**, 119 (2009).
161. ZEUS Collab.: M. Derrick *et al.*, Z. Phys. **C68**, 29 (1995); ZEUS Collab.: J. Breitweg *et al.*, Eur. Phys. J. **C2**, 77 (1998).
162. ZEUS Collab.: S. Chekanov *et al.*, Phys. Lett. **B553**, 141 (2003).
163. ZEUS Collab.: S. Chekanov *et al.*, Nucl. Phys. **B786**, 181 (2007).
164. H1 Collab.: F. D. Aaron *et al.*, Eur. Phys. J. **C61**, 185 (2009).
165. H1 Collab.: C. Adloff *et al.*, Eur. Phys. J. **C18**, 293 (2000); H1 Collab.: A. Aktas *et al.*, Eur. Phys. J. **C36**, 413 (2004).
166. P. Dixon *et al.*, J. Phys. **G25**, 1453 (1999).
167. H1 Collab.: C. Adloff *et al.*, Phys. Lett. **B462**, 440 (1999).
168. D. Graudenz, Fortsch. Phys. **45**, 629 (1997).
169. S. Albino *et al.*, Phys. Rev. **D75**, 034018 (2007).
170. OPAL Collab.: P.D. Acton *et al.*, Phys. Lett. **B305**, 407 (1993).
171. E632 Collab.: D. DeProspo *et al.*, Phys. Rev. **D50**, 6691 (1994).
172. ZEUS Collab.: S. Chekanov *et al.*, Eur. Phys. J. **C51**, 1 (2007).
173. G. Grindhammer *et al.*, in: *Proceedings of the Workshop on Monte Carlo Generators for HERA Physics*, Hamburg, Germany, 1998/1999.
174. N. Brook *et al.*, in: *Proceedings of the Workshop for Future HERA Physics at HERA*, Hamburg, Germany, 1996.
175. CDF Collab.: F. Abe *et al.*, Phys. Rev. Lett. **61**, 1819 (1988).
176. CDF Collab.: D.E. Acosta *et al.*, Phys. Rev. **D72**, 052001 (2005).
177. UA1 Collab.: G. Arnison *et al.*, Phys. Lett. **B118**, 167 (1982).
178. UA1 Collab.: C. Albajar *et al.*, Nucl. Phys. **B335**, 261 (1990).
179. UA1 Collab.: G. Bocquet *et al.*, Phys. Lett. **B366**, 434 (1996).
180. UA2 Collab.: M. Banner *et al.*, Phys. Lett. **B122**, 322 (1983).
181. UA2 Collab.: M. Banner *et al.*, Phys. Lett. **B115**, 59 (1982).
182. UA2 Collab.: M. Banner *et al.*, Z. Phys. **C27**, 329 (1985).
183. PHENIX Collab.: S. S. Adler *et al.*, Phys. Rev. Lett. **91**, 241803 (2003).
184. PHENIX Collab.: A. Adare *et al.*, Phys. Rev. **D76**, 051106 (2007).
185. BRAHMS Collab.: I. Arsene *et al.*, Phys. Rev. Lett. **98**, 252001 (2007).
186. STAR Collab.: J. Adams *et al.*, Phys. Lett. **B637**, 161 (2006); STAR Collab.: J. Adams *et al.*, Phys. Rev. Lett. **97**, 152302 (2006); STAR Collab.: B.I. Abelev *et al.*, Phys. Rev. **C75**, 064901 (2007); STAR Collab.: G. Agakishiev *et al.*, Phys. Rev. Lett. **108**, 072302 (2012); STAR Collab.: B.I. Abelev *et al.*, Phys. Rev. **C81**, 064904 (2010).
187. ALICE Collab.: B. Abelev *et al.*, Phys. Lett. **B717**, 162 (2012).
188. ALICE Collab.: B. Abelev *et al.*, Eur. Phys. J. **C73**, 2662 (2013).
189. E706 Collab.: L. Apanasevich *et al.*, Phys. Rev. Lett. **81**, 2642 (1998).
190. UA6 Collab.: G. Ballocci *et al.*, Phys. Lett. **B436**, 222 (1998).
191. WA70 Collab.: M. Bonesini *et al.*, Z. Phys. **C38**, 371 (1988).
192. AFS Collab.: E. Anassontzis *et al.*, Sov. J. Nucl. Phys. **51**, 836 (1990).
193. R806 Collab.: C. Kourkouvelis *et al.*, Z. Phys. **C5**, 95 (1980).
194. E706 Collab.: L. Apanasevich *et al.*, Phys. Rev. **D68**, 052001 (2003).
195. ALICE Collab.: K. Aamodt *et al.*, Eur. Phys. J. **C71**, 1594 (2011); ALICE Collab.: B. Abelev *et al.*, Phys. Lett. **B710**, 557 (2012); ALICE Collab.: B. Abelev *et al.*, Phys. Lett. **B712**, 309 (2012); ALICE Collab.: B. Abelev *et al.*, Eur. Phys. J. **C75**, 1 (2015); ALICE Collab.: J. Adam *et al.*, Eur. Phys. J. **C75**, 226 (2015).

196. ATLAS Collab.: G. Aad *et al.*, Phys. Rev. **D85**, 012001 (2012); ATLAS Collab.: G. Aad *et al.*, Eur. Phys. J. **C74**, 2895 (2014).
197. CDF Collab.: D. Acosta *et al.*, Phys. Rev. **D72**, 052001 (2005).
198. CMS Collab.: V. Khachatryan *et al.*, JHEP **1105**, 064 (2011); CMS Collab.: S. Chatrchyan *et al.*, Eur. Phys. J. **C72**, 2164 (2012); CMS Collab.: S. Chatrchyan *et al.*, Phys. Rev. **D88**, 052001 (2013).
199. LHCb Collab.: R. Aaij *et al.*, Phys. Lett. **B703**, 267 (2011).
200. F. Aversa *et al.*, Nucl. Phys. **B327**, 105 (1989); D. de Florian, Phys. Rev. **D67**, 054004 (2003); B. Jager *et al.*, Phys. Rev. **D67**, 054005 (2003).
201. U. Baur *et al.*, hep-ph/0005226 (2000).
202. P. Aurenche, *et al.*, Eur. Phys. J. **C13**, 347 (2000).
203. L. Apanasevich *et al.*, Phys. Rev. **D59**, 074007 (1999).
204. U. D'Alesio and F. Murgia, Phys. Rev. **D70**, 074009 (2004).
205. D. de Florian and W. Vogelsang, Phys. Rev. **D71**, 114004 (2005).
206. L.G. Almeida *et al.*, Phys. Rev. **D80**, 074016 (2009).
207. D. d'Enterria *et al.*, Nucl. Phys. **B883**, 615 (2014).
208. PHENIX Collab.: A. Adare *et al.*, Phys. Rev. **D76**, 051106 (2007).
209. PHENIX Collab.: A. Adare *et al.*, Phys. Rev. **D79**, 012003 (2009).
210. PHENIX Collab.: K. Adcox *et al.*, Phys. Rev. Lett. **88**, 022301 (2002); STAR Collab.: C. Adler *et al.*, Phys. Rev. Lett. **90**, 082302 (2003).
211. COMPASS Collab.: M. Alekseev *et al.*, Phys. Lett. **B660**, 458, (2008).
212. HERMES Collab.: A. Airapetian *et al.*, Phys. Rev. **D71**, 012003 (2005).
213. SMC Collab.: B. Adeva *et al.*, Phys. Lett. **B420**, 180 (1998).
214. HERMES Collab.: A. Airapetian *et al.*, Phys. Lett. **B666**, 446 (2008).
215. D. de Florian *et al.*, Phys. Rev. Lett. **101**, 072001 (2008).
216. P.J. Mulders and R.D. Tangerman, Nucl. Phys. **B461**, 197 (1996); Erratum *ibid.*, **B484**, 538 (1997).
217. R. Jacob, Nucl. Phys. **A711**, 35 (2002).
218. COMPASS Collab.: M. Alekseev *et al.*, Eur. Phys. J. **C64**, 171 (2009).
219. HERMES Collab.: A. Airapetian *et al.*, Phys. Rev. **D74**, 072004 (2006).
220. J.P. Ralston and D.E. Soper, Nucl. Phys. **B152**, 109 (1979).
221. J. Collins, Nucl. Phys. **B396**, 161 (1993).
222. D. Sivers, Phys. Rev. **D43**, 261 (1991).
223. CLAS Collab.: H. Avakian *et al.*, Phys. Rev. **D69**, 112004 (2004).
224. HERMES Collab.: A. Airapetian *et al.*, Phys. Rev. Lett. **84**, 4047 (2000).
225. HERMES Collab.: A. Airapetian *et al.*, Phys. Rev. **D64**, 097101 (2001).
226. HERMES Collab.: A. Airapetian *et al.*, Phys. Rev. Lett. **94**, 012002 (2005).
227. BELLE Collab.: K. Abe *et al.*, Phys. Rev. Lett. **96**, 232002 (2006).
228. BELLE Collab.: K. Abe *et al.*, Phys. Rev. **D78**, 032011 (2008).
229. COMPASS Collab.: V.Y. Alexakhin *et al.*, Phys. Rev. Lett. **94**, 202002 (2005).
230. COMPASS Collab.: V.Y. Alexakhin *et al.*, Nucl. Phys. **B765**, 31 (2007).
231. COMPASS Collab.: M. Alekseev *et al.*, Phys. Lett. **B673**, 127 (2009).
232. COMPASS Collab.: M. Alekseev *et al.*, Phys. Lett. **B692**, 240 (2010).
233. COMPASS Collab.: M. Alekseev *et al.*, Eur. Phys. J. **C70**, 39 (2010).
234. V.A. Khoze *et al.*, *Proceedings, Conference on High-Energy Physics, Tbilisi 1976*; J.D. Bjorken, Phys. Rev. **D17**, 171 (1978).
235. B. Mele and P. Nason, Phys. Lett. **B245**, 635 (1990); B. Mele and P. Nason, Nucl. Phys. **B361**, 626 (1991).
236. P. Nason and C. Oleari, Phys. Lett. **B418**, 199 (1998); P. Nason and C. Oleari, Phys. Lett. **B447**, 327 (1999); P. Nason and C. Oleari, Nucl. Phys. **B565**, 245 (2000).
237. K. Melnikov and A. Mitov, Phys. Rev. **D70**, 034027 (2004).
238. C. Peterson *et al.*, Phys. Rev. **D27**, 105 (1983).
239. V.G. Kartvelishvili *et al.*, Phys. Lett. **B78**, 615 (1978).
240. P. Collins and T. Spiller, J. Phys. **G11**, 1289 (1985).
241. G. Colangelo and P. Nason, Phys. Lett. **B285**, 167 (1992).
242. M.G. Bowler, Z. Phys. **C11**, 169 (1981).
243. E. Braaten *et al.*, Phys. Rev. **D51**, 4819 (1995).
244. OPAL Collab.: R. Akers *et al.*, Z. Phys. **C67**, 27 (1995).
245. Particle Data Group: C. Amsler *et al.*, Phys. Lett. **B667**, 1 (2008).
246. J. Chrin, Z. Phys. **C36**, 163 (1987).
247. R.L. Jaffe and L. Randall, Nucl. Phys. **B412**, 79 (1994).
248. M. Cacciari and E. Gardi, Nucl. Phys. **B664**, 299 (2003).
249. L. Randall and N. Rius, Nucl. Phys. **B441**, 167 (1995).
250. J. Collins, Phys. Rev. **D58**, 094002 (1998).
251. B.A. Kniehl *et al.*, Eur. Phys. J. **C41**, 199 (2005).
252. CLEO Collab.: M. Artuso *et al.*, Phys. Rev. **D70**, 112001 (2004).
253. BELLE Collab.: R. Seuster *et al.*, Phys. Rev. **D73**, 032002 (2006).
254. CLEO Collab.: R.A. Briere *et al.*, Phys. Rev. **D62**, 112003 (2000).
255. BABAR Collab.: B. Aubert *et al.*, Phys. Rev. **D65**, 091104 (2002).
256. CLEO Collab.: D. Bortoletto *et al.*, Phys. Rev. **D37**, 1719 (1988).
257. ARGUS Collab.: H. Albrecht *et al.*, Z. Phys. **C52**, 353 (1991).
258. ARGUS Collab.: H. Albrecht *et al.*, Z. Phys. **C54**, 1 (1992).
259. ALEPH Collab.: R. Barate *et al.*, Phys. Lett. **B561**, 213 (2003).
260. H1 Collab.: F.D. Aaron *et al.*, Eur. Phys. J. **C65**, 89 (2010).
261. ZEUS Collab.: S. Chekanov *et al.*, JHEP **0707**, 074 (2007); ZEUS Collab.: H. Abramowicz *et al.*, JHEP **1309** (2013); H1 Collab.: A. Aktas *et al.*, Eur. Phys. J. **C51**, 271 (2007); H1 Collab.: F.D. Aaron *et al.*, Eur. Phys. J. **C59**, 589 (2009).
262. ALEPH Collab.: D. Buskulic *et al.*, Phys. Lett. **B357**, 699 (1995).
263. ALEPH Collab.: A. Heister *et al.*, Phys. Lett. **B512**, 30 (2001); DELPHI Collab.: J. Abdallah *et al.*, Eur. Phys. J. **C71**, 1557 (2011); OPAL Collab.: G. Abbiendi *et al.*, Eur. Phys. J. **C29**, 463 (2003); SLD Collab.: K. Abe *et al.*, Phys. Rev. **D65**, 092006 (2002); Erratum *ibid.*, **D66**, 079905 (2002).
264. L3 Collab.: B. Adeva *et al.*, Phys. Lett. **B261**, 177 (1991).
265. CDF Collab.: F. Abe *et al.*, Phys. Rev. Lett. **71**, 500 (1993); CDF Collab.: F. Abe *et al.*, Phys. Rev. Lett. **71**, 2396 (1993); CDF Collab.: F. Abe *et al.*, Phys. Rev. **D50**, 4252 (1994); CDF Collab.: F. Abe *et al.*, Phys. Rev. Lett. **75**, 1451 (1995); CDF Collab.: D. Acosta *et al.*, Phys. Rev. **D66**, 032002 (2002); CDF Collab.: D. Acosta *et al.*, Phys. Rev. **D65**, 052005 (2002); D0 Collab.: S. Abachi *et al.*, Phys. Rev. Lett. **74**, 3548 (1995); UA1 Collab.: C. Albajar *et al.*, Phys. Lett. **B186**, 237 (1987); UA1 Collab.: C. Albajar *et al.*, Phys. Lett. **B256**, 121 (1991); Erratum *ibid.*, **B272**, 497 (1991).
266. O. Biebel, P. Nason, and B.R. Webber, Bicecca-FT-01-20, Cavendish-HEP-01/12, MPI-PhE/2001-14 [hep-ph/0109282 (2001)].
267. H. Evans, arXiv:1110.5294 (2011); E. Aguiló, arXiv:1205.5678 (2012); F. Simonetto, Journal of Physics: Conference Series **347**, 012014 (2012).
268. CMS Collab.: V. Khachatryan *et al.*, JHEP **1103**, 136 (2011); ATLAS Collab.: G. Aad *et al.*, Eur. Phys. J. **C71**, 1846 (2011); CMS Collab.: S. Chatrchyan *et al.*, JHEP **1204**, 084 (2012); ATLAS Collab.: G. Aad *et al.*, Eur. Phys. J. **C73**, 2301 (2013).
269. H. Jung *et al.*, Phys. Rev. **D85**, 034035 (2012).

270. CDF Collab.: D. Acosta *et al.*, Phys. Rev. Lett. **91**, 241804 (2003);  
CDF Collab.: D. Acosta *et al.*, Phys. Rev. **D71**, 032001 (2005).
271. M. Cacciari and P. Nason, JHEP **0309**, 006 (2003);  
M. Cacciari *et al.*, JHEP **0407**, 033 (2004);  
B.A. Kniehl *et al.*, Phys. Rev. Lett. **96**, 012001 (2006).
272. ZEUS Collab.: H. Abramovitz *et al.*, Eur. Phys. J. **C71**, 1573 (2011).
273. ZEUS Collab.: S. Chekanov *et al.*, Phys. Rev. **D78**, 072001 (2008).
274. ZEUS Collab.: S. Chekanov *et al.*, JHEP **0902**, 032 (2009).
275. H1 Collab.: F.D. Aaron *et al.*, Eur. Phys. J. **C72**, 2148 (2012).
276. A.H. Mueller and P. Nason, Nucl. Phys. **B266**, 265 (1986);  
M.L. Mangano and P. Nason, Phys. Lett. **B285**, 160 (1992).
277. M.H. Seymour, Nucl. Phys. **B436**, 163 (1995).
278. D.J. Miller and M.H. Seymour, Phys. Lett. **B435**, 213 (1998).
279. ALEPH Collab.: R. Barate *et al.*, Phys. Lett. **B434**, 437 (1998).
280. DELPHI Collab.: P. Abreu *et al.*, Phys. Lett. **B405**, 202 (1997).
281. L3 Collab.: M. Acciarri *et al.*, Phys. Lett. **B476**, 243 (2000).
282. OPAL Collab.: G. Abbiendi *et al.*, Eur. Phys. J. **C13**, 1 (2000).
283. SLD Collab.: K. Abe *et al.*, SLAC-PUB-8157 [**hep-ex/9908028** (1999)].

## 21. EXPERIMENTAL TESTS OF GRAVITATIONAL THEORY

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Einstein's General Relativity, the current “standard” theory of gravitation, describes gravity as a universal deformation of the Minkowski metric:

$$g_{\mu\nu}(x^\lambda) = \eta_{\mu\nu} + h_{\mu\nu}(x^\lambda), \text{ where } \eta_{\mu\nu} = \text{diag}(-1, +1, +1, +1). \quad (21.1)$$

General Relativity is classically defined by two postulates. One postulate states that the Lagrangian density describing the propagation and self-interaction of the gravitational field is

$$\mathcal{L}_{\text{Ein}}[g_{\alpha\beta}] = \frac{c^4}{16\pi G_N} \sqrt{g} g^{\mu\nu} R_{\mu\nu}(g_{\alpha\beta}), \quad (21.2)$$

where  $G_N$  is Newton's constant,  $g = -\det(g_{\mu\nu})$ ,  $g^{\mu\nu}$  is the matrix inverse of  $g_{\mu\nu}$ , and where the Ricci tensor  $R_{\mu\nu} \equiv R^\alpha_{\mu\alpha\nu}$  is the only independent trace of the curvature tensor

$$R^\alpha_{\mu\beta\nu} = \partial_\beta \Gamma^\alpha_{\mu\nu} - \partial_\nu \Gamma^\alpha_{\mu\beta} + \Gamma^\alpha_{\sigma\beta} \Gamma^\sigma_{\mu\nu} - \Gamma^\alpha_{\sigma\nu} \Gamma^\sigma_{\mu\beta}, \quad (21.3)$$

$$\Gamma^\lambda_{\mu\nu} = \frac{1}{2} g^{\lambda\sigma} (\partial_\mu g_{\nu\sigma} + \partial_\nu g_{\mu\sigma} - \partial_\sigma g_{\mu\nu}), \quad (21.4)$$

A second postulate states that  $g_{\mu\nu}$  couples universally, and minimally, to all the fields of the Standard Model by replacing everywhere the Minkowski metric  $\eta_{\mu\nu}$ . Schematically (suppressing matrix indices and labels for the various gauge fields and fermions and for the Higgs doublet),

$$\mathcal{L}_{\text{SM}}[\psi, A_\mu, H, g_{\mu\nu}] = -\frac{1}{4} \sum \sqrt{g} g^{\mu\alpha} g^{\nu\beta} F_{\mu\nu}^a F_{\alpha\beta}^a - \sum \sqrt{g} \bar{\psi} \gamma^\mu D_\mu \psi - \frac{1}{2} \sqrt{g} g^{\mu\nu} \bar{D}_\mu H D_\nu H - \sqrt{g} V(H) - \sum \lambda \sqrt{g} \bar{\psi} H \psi, \quad (21.5)$$

where  $\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2g^{\mu\nu}$ , and where the covariant derivative  $D_\mu$  contains, besides the usual gauge field terms, a spin-dependent gravitational contribution. From the total action follow Einstein's field equations,

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \frac{8\pi G_N}{c^4} T_{\mu\nu}. \quad (21.6)$$

Here  $R = g^{\mu\nu} R_{\mu\nu}$ ,  $T_{\mu\nu} = g_{\mu\alpha} g_{\nu\beta} T^{\alpha\beta}$ , and  $T^{\mu\nu} = (2/\sqrt{g}) \delta \mathcal{L}_{\text{SM}} / \delta g_{\mu\nu}$  is the (symmetric) energy-momentum tensor of the Standard Model matter. The theory is invariant under arbitrary coordinate transformations:  $x'^\mu = f^\mu(x^\nu)$ . To solve the field equations Eq. (21.6), one needs to fix this coordinate gauge freedom. *E.g.*, the “harmonic gauge” (which is the analogue of the Lorenz gauge,  $\partial_\mu A^\mu = 0$ , in electromagnetism) corresponds to imposing the condition  $\partial_\nu (\sqrt{g} g^{\mu\nu}) = 0$ .

In this *Review*, we only consider the classical limit of gravitation (*i.e.* classical matter and classical gravity). Quantum gravitational effects are expected (when considered at low energy) to correct the classical action Eq. (21.3) by additional terms involving quadratic and higher powers of the curvature tensor. This suggests that the validity of classical gravity extends (at most) down to length scales of order the Planck length  $L_P = \sqrt{\hbar G_N / c^3} \simeq 1.62 \times 10^{-33}$  cm, *i.e.* up to energy scales of order the Planck energy  $E_P = \sqrt{\hbar c^5 / G_N} \simeq 1.22 \times 10^{19}$  GeV. Considering quantum matter in a classical gravitational background also poses interesting challenges, notably the possibility that the zero-point fluctuations of the matter fields generate a nonvanishing vacuum energy density  $\rho_{\text{vac}}$ , corresponding to a term  $-\sqrt{g} \rho_{\text{vac}}$  in  $\mathcal{L}_{\text{SM}}$  [1]. This is equivalent to adding a “cosmological constant” term  $+\Lambda g_{\mu\nu}$  on the left-hand side of Einstein's equations Eq. (21.6), with  $\Lambda = 8\pi G_N \rho_{\text{vac}} / c^4$ . Recent cosmological observations (see the following *Reviews*) suggest a positive value of  $\Lambda$  corresponding to  $\rho_{\text{vac}} \approx (2.3 \times 10^{-3} \text{ eV})^4$ . Such a small value has a negligible effect on the non cosmological tests discussed below.

### 21.1. Experimental tests of the coupling between matter and gravity

The universality of the coupling between  $g_{\mu\nu}$  and the Standard Model matter postulated in Eq. (21.5) (“Equivalence Principle”) has many observable consequences [2]. First, it predicts that the outcome of a local non-gravitational experiment, referred to local standards, does not depend on where, when, and in which locally inertial frame, the experiment is performed. This means, for instance, that local experiments should neither feel the cosmological evolution of the universe (constancy of the “constants”), nor exhibit preferred directions in spacetime (isotropy of space, local Lorentz invariance). These predictions are consistent with many experiments and observations. Stringent limits on a possible time variation of the basic coupling constants have been obtained by analyzing a natural fission reactor phenomenon which took place at Oklo, Gabon, two billion years ago [3,4]. These limits are at the  $1 \times 10^{-7}$  level for the fractional variation of the fine-structure constant  $\alpha_{\text{em}}$  [4], and at the  $4 \times 10^{-9}$  level for the fractional variation of the ratio  $m_q / \Lambda_{\text{QCD}}$  between the light quark masses and  $\Lambda_{\text{QCD}}$  [5]. The determination of the lifetime of Rhenium 187 from isotopic measurements of some meteorites dating back to the formation of the solar system (about 4.6 Gyr ago) yields comparably strong limits [6]. Measurements of absorption lines in astronomical spectra also give stringent limits on the variability of both  $\alpha_{\text{em}}$  and  $\mu = m_p / m_e$  at cosmological redshifts. *E.g.*

$$\Delta \alpha_{\text{em}} / \alpha_{\text{em}} = (1.3 \pm 2.4_{\text{stat}} \pm 1.0_{\text{sys}}) \times 10^{-6} \quad (21.7)$$

at a redshift  $z = 1.6919$  [7], and

$$|\Delta \mu / \mu| < 4 \times 10^{-7} (95\% \text{ C.L.}), \quad (21.8)$$

at a redshift  $z = 0.88582$  [8]. There are also strong limits (at the  $10^{-5}$  level) on the variation of  $\alpha_{\text{em}}$  and  $\mu = m_p / m_e$  at very large redshifts, such as  $z = 4.22$  [9] and  $z = 5.2$  [10]. Direct laboratory limits (based on monitoring the frequency ratio of several different atomic clocks) on the present time variation of  $\alpha_{\text{em}}$ ,  $\mu = m_p / m_e$ , and  $m_q / \Lambda_{\text{QCD}}$  have reached the levels [11]:

$$d \ln(\alpha_{\text{em}}) / dt = (-2.5 \pm 2.6) \times 10^{-17} \text{ yr}^{-1},$$

$$d \ln(\mu) / dt = (-1.5 \pm 3.0) \times 10^{-16} \text{ yr}^{-1},$$

$$d \ln(m_q / \Lambda_{\text{QCD}}) / dt = (7.1 \pm 4.4) \times 10^{-15} \text{ yr}^{-1}. \quad (21.9)$$

There are also experimental limits on a possible dependence of coupling constants on the gravitational potential [11,12]. See Ref. 13 for a review of the issue of “variable constants.”

The highest precision tests of the isotropy of space have been performed by looking for possible quadrupolar shifts of nuclear energy levels [14]. The (null) results can be interpreted as testing the fact that the various pieces in the matter Lagrangian Eq. (21.5) are indeed coupled to one and the same external metric  $g_{\mu\nu}$  to the  $10^{-29}$  level. For astrophysical constraints on possible Planck-scale violations of Lorentz invariance, see Ref. 15.

The universal coupling to  $g_{\mu\nu}$  postulated in Eq. (21.5) implies that two (electrically neutral) test bodies dropped at the same location and with the same velocity in an external gravitational field fall in the same way, independently of their masses and compositions. The universality of the acceleration of free fall has been verified at the  $10^{-13}$  level for laboratory bodies, notably Beryllium-Titanium, and Beryllium-Aluminum test bodies [16,17],

$$(\Delta a / a)_{\text{BeTi}} = (0.3 \pm 1.8) \times 10^{-13},$$

$$(\Delta a / a)_{\text{BeAl}} = (-0.7 \pm 1.3) \times 10^{-13}. \quad (21.10)$$

The universality of free fall has also been verified when comparing the fall of classical and quantum objects ( $6 \times 10^{-9}$  level [18]), or of two quantum objects ( $5 \times 10^{-7}$  level [19]). The gravitational accelerations of the Earth and the Moon toward the Sun have also been verified to agree [20],

$$(\Delta a / a)_{\text{EarthMoon}} = (-0.8 \pm 1.3) \times 10^{-13}. \quad (21.11)$$

The latter result constrains not only how  $g_{\mu\nu}$  couples to matter, but also how it couples to itself [21] (“strong equivalence principle”). See also Ref. 22 for a review of torsion balance experiments.

Finally, Eq. (21.5) also implies that two identically constructed clocks located at two different positions in a static external Newtonian potential  $U(\mathbf{x}) = \sum G_N m/r$  exhibit, when intercompared by means of electromagnetic signals, the (apparent) difference in clock rate,  $\tau_1/\tau_2 = \nu_2/\nu_1 = 1 + [U(\mathbf{x}_1) - U(\mathbf{x}_2)]/c^2 + O(1/c^4)$ , independently of their nature and constitution. This universal gravitational redshift of clock rates has been verified at the  $10^{-4}$  level by comparing a hydrogen-maser clock flying on a rocket up to an altitude  $\sim 10,000$  km to a similar clock on the ground [23]. The redshift due to a height change of only 33 cm has been detected by comparing two optical clocks based on  $^{27}\text{Al}^+$  ions [24].

## 21.2. Tests of the dynamics of the gravitational field in the weak field regime

The effect on matter of one-graviton exchange, *i.e.*, the interaction Lagrangian obtained when solving Einstein’s field equations Eq. (21.6) written in, say, the harmonic gauge at first order in  $h_{\mu\nu}$ ,

$$\square h_{\mu\nu} = -\frac{16\pi G_N}{c^4} (T_{\mu\nu} - \frac{1}{2} T \eta_{\mu\nu}) + O(h^2) + O(hT), \quad (21.12)$$

reads  $-(8\pi G_N/c^4) T^{\mu\nu} \square^{-1} (T_{\mu\nu} - \frac{1}{2} T \eta_{\mu\nu})$ . For a system of  $N$  moving point masses, with free Lagrangian  $L^{(1)} = \sum_{A=1}^N -m_A c^2 \sqrt{1 - v_A^2/c^2}$ ,

this interaction, expanded to order  $v^2/c^2$ , reads (with  $r_{AB} \equiv |\mathbf{x}_A - \mathbf{x}_B|$ ,  $\mathbf{n}_{AB} \equiv (\mathbf{x}_A - \mathbf{x}_B)/r_{AB}$ )

$$L^{(2)} = \frac{1}{2} \sum_{A \neq B} \frac{G_N m_A m_B}{r_{AB}} \left[ 1 + \frac{3}{2c^2} (v_A^2 + v_B^2) - \frac{7}{2c^2} (\mathbf{v}_A \cdot \mathbf{v}_B) - \frac{1}{2c^2} (\mathbf{n}_{AB} \cdot \mathbf{v}_A) (\mathbf{n}_{AB} \cdot \mathbf{v}_B) + O\left(\frac{1}{c^4}\right) \right]. \quad (21.13)$$

The two-body interactions, Eq. (21.13), exhibit  $v^2/c^2$  corrections to Newton’s  $1/r$  potential induced by spin-2 exchange (“gravito-magnetism”). Consistency at the “post-Newtonian” level  $v^2/c^2 \sim G_N m/r c^2$  requires that one also considers the three-body interactions induced by some of the three-graviton vertices and other nonlinearities (terms  $O(h^2)$  and  $O(hT)$  in Eq. (21.12)),

$$L^{(3)} = -\frac{1}{2} \sum_{B \neq A \neq C} \frac{G_N^2 m_A m_B m_C}{r_{AB} r_{AC} c^2} + O\left(\frac{1}{c^4}\right). \quad (21.14)$$

All currently performed gravitational experiments in the solar system, including perihelion advances of planetary orbits, the bending and delay of electromagnetic signals passing near the Sun, and very accurate ranging data to the Moon obtained by laser echoes, are compatible with the post-Newtonian results Eqs. (21.12)–(21.14). The “gravito-magnetic” interactions  $\propto \mathbf{v}_A \mathbf{v}_B$  contained in Eq. (21.13) are involved in many of these experimental tests. They have been particularly tested in lunar laser ranging data [20], in the LAGEOS satellite observations [25,26], and in the dedicated Gravity Probe B mission [27]. The recently launched LARES satellite promises to improve the accuracy of such tests [26].

Similar to what is done in discussions of precision electroweak experiments, it is useful to quantify the significance of precision gravitational experiments by parameterizing plausible deviations from General Relativity. Here, we shall focus on the simplest, and most conservative deviations from Einstein’s pure spin-2 theory defined by adding new, bosonic light or massless, macroscopically coupled fields. [For discussions of less conservative deviations from Einstein’s theory (modified newtonian dynamics, massive gravity, higher-order gravity,  $f(R)$ -gravity, Lorentz-violating theories,...) and their confrontation with experiment, see [28,29,30].] The possibility of new gravitational-strength couplings leading (on small, and possibly large, scales) to

deviations from Einsteinian (and Newtonian) gravity is suggested by String Theory [31], and by Brane World ideas [32]. For reviews of experimental constraints on Yukawa-type additional interactions, see Refs. [22,33,17]. Experiments have set limits on non-Newtonian forces down to 0.056 mm [34].

Here, we shall focus on the parametrization of long-range deviations from relativistic gravity obtained by adding a strictly massless (*i.e.* without self-interaction  $V(\varphi) = 0$ ) scalar field  $\varphi$  coupled to the trace of the energy-momentum tensor  $T = g_{\mu\nu} T^{\mu\nu}$  [35]. The most general such theory contains an arbitrary function  $a(\varphi)$  of the scalar field, and can be defined by the Lagrangian

$$\mathcal{L}_{\text{tot}}[g_{\mu\nu}, \varphi, \psi, A_\mu, H] = \frac{c^4}{16\pi G} \sqrt{g} (R(g_{\mu\nu}) - 2g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi) + \mathcal{L}_{\text{SM}}[\psi, A_\mu, H, \tilde{g}_{\mu\nu}], \quad (21.15)$$

where  $G$  is a “bare” Newton constant, and where the Standard Model matter is coupled not to the “Einstein” (pure spin-2) metric  $g_{\mu\nu}$ , but to the conformally related (“Jordan-Fierz”) metric  $\tilde{g}_{\mu\nu} = \exp(2a(\varphi))g_{\mu\nu}$ . The scalar field equation  $\square_g \varphi = -(4\pi G/c^4) \alpha(\varphi) T$  displays  $\alpha(\varphi) \equiv \partial a(\varphi)/\partial \varphi$  as the basic (field-dependent) coupling between  $\varphi$  and matter [36]. The one-parameter ( $\omega$ ) Jordan-Fierz-Brans-Dicke theory [35] is the special case  $a(\varphi) = \alpha_0 \varphi$  leading to a field-independent coupling  $\alpha(\varphi) = \alpha_0$  (with  $\alpha_0^2 = 1/(2\omega + 3)$ ). The addition of a self-interaction term  $V(\varphi)$  in Eq. (21.15) introduces new phenomenological possibilities; notably the “chameleon mechanism” [37].

In the weak-field slow-motion limit appropriate to describing gravitational experiments in the solar system, the addition of  $\varphi$  modifies Einstein’s predictions only through the appearance of two “post-Einstein” dimensionless parameters:  $\bar{\gamma} = -2\alpha_0^2/(1 + \alpha_0^2)$  and  $\bar{\beta} = +\frac{1}{2}\beta_0\alpha_0^2/(1 + \alpha_0^2)^2$ , where  $\alpha_0 \equiv \alpha(\varphi_0)$ ,  $\beta_0 \equiv \partial \alpha(\varphi_0)/\partial \varphi_0$ ,  $\varphi_0$  denoting the vacuum expectation value of  $\varphi$ . These parameters show up also naturally (in the form  $\gamma_{\text{PPN}} = 1 + \bar{\gamma}$ ,  $\beta_{\text{PPN}} = 1 + \bar{\beta}$ ) in phenomenological discussions of possible deviations from General Relativity [2]. The parameter  $\bar{\gamma}$  measures the admixture of spin 0 to Einstein’s graviton, and contributes an extra term  $+\bar{\gamma}(\mathbf{v}_A - \mathbf{v}_B)^2/c^2$  in the square brackets of the two-body Lagrangian Eq. (21.13). The parameter  $\bar{\beta}$  modifies the three-body interaction Eq. (21.14) by an overall multiplicative factor  $1 + 2\bar{\beta}$ . Moreover, the combination  $\eta \equiv 4\bar{\beta} - \bar{\gamma}$  parameterizes the lowest order effect of the self-gravity of orbiting masses by modifying the Newtonian interaction energy terms in Eq. (21.13) into  $G_{AB} m_A m_B / r_{AB}$ , with a body-dependent gravitational “constant”  $G_{AB} = G_N [1 + \eta (E_A^{\text{grav}}/m_A c^2 + E_B^{\text{grav}}/m_B c^2) + O(1/c^4)]$ , where  $G_N = G \exp[2a(\varphi_0)](1 + \alpha_0^2)$  and where  $E_A^{\text{grav}}$  denotes the gravitational binding energy of body  $A$ .

The best current limits on the post-Einstein parameters  $\bar{\gamma}$  and  $\bar{\beta}$  are (at the 68% confidence level):

$$\bar{\gamma} = (2.1 \pm 2.3) \times 10^{-5}, \quad (21.16)$$

deduced from the additional Doppler shift experienced by radio-wave beams connecting the Earth to the Cassini spacecraft when they passed near the Sun [38], and

$$|\bar{\beta}| < 7 \times 10^{-5}, \quad (21.17)$$

from a study of the global sensitivity of planetary ephemerides to post-Einstein parameters [39]. More stringent limits on  $\bar{\gamma}$  are obtained in models (*e.g.*, string-inspired ones [31]) where scalar couplings violate the Equivalence Principle.

## 21.3. Tests of the dynamics of the gravitational field in the radiative and/or strong field regimes

The discovery of pulsars (*i.e.*, rotating neutron stars emitting a beam of radio noise) in gravitationally bound orbits [40,41] has opened up an entirely new testing ground for relativistic gravity, giving us an experimental handle on the regime of radiative and/or strong gravitational fields. In these systems, the finite velocity of propagation

of the gravitational interaction between the pulsar and its companion generates damping-like terms at order  $(v/c)^5$  in the equations of motion [43]. These damping forces are the local counterparts of the gravitational radiation emitted at infinity by the system (“gravitational radiation reaction”). They cause the binary orbit to shrink and its orbital period  $P_b$  to decrease. The remarkable stability of pulsar clocks has allowed one to measure the corresponding very small orbital period decay  $\dot{P}_b \equiv dP_b/dt \sim -(v/c)^5 \sim -10^{-12}$  in several binary systems, thereby giving us a direct experimental confirmation of the propagation properties of the gravitational field, and, in particular, an experimental confirmation that the speed of propagation of gravity  $c_g$  is equal to the velocity of light  $c$  to better than a part in a thousand. [See, also, [42] for tight constraints on the difference  $c - c_g$ , when it is assumed positive.] In addition, the surface gravitational potential of a neutron star  $h_{00}(R_{\text{NS}}) \simeq 2Gm/c^2 R_{\text{NS}} \simeq 0.4$  being a factor  $\sim 10^8$  higher than the surface potential of the Earth, and a mere factor 2.5 below the black hole limit ( $h_{00}(R_{\text{BH}}) = 1$ ), pulsar data have allowed one to obtain several accurate tests of the strong-gravitational-field regime, as we discuss next.

Binary pulsar timing data record the times of arrival of successive electromagnetic pulses emitted by a pulsar orbiting around the center of mass of a binary system. After correcting for the Earth motion around the Sun and for the dispersion due to propagation in the interstellar plasma, the time of arrival of the  $N$ th pulse  $t_N$  can be described by a generic, parameterized “timing formula” [44] whose functional form is common to the whole class of tensor-scalar gravitation theories:

$$t_N - t_0 = F[T_N(\nu_p, \dot{\nu}_p, \ddot{\nu}_p); \{p^K\}; \{p^{PK}\}] . \quad (21.18)$$

Here,  $T_N$  is the pulsar proper time corresponding to the  $N$ th turn given by  $N/2\pi = \nu_p T_N + \frac{1}{2}\dot{\nu}_p T_N^2 + \frac{1}{6}\ddot{\nu}_p T_N^3$  (with  $\nu_p \equiv 1/P_p$  the spin frequency of the pulsar, *etc.*),  $\{p^K\} = \{P_b, T_0, e, \omega_0, x\}$  is the set of “Keplerian” parameters (notably, orbital period  $P_b$ , eccentricity  $e$ , periastron longitude  $\omega_0$  and projected semi-major axis  $x = a \sin i/c$ ), and  $\{p^{PK}\} = \{k, \gamma_{\text{timing}}, \dot{P}_b, r, s, \delta_\theta, \dot{e}, \dot{x}\}$  denotes the set of (separately measurable) “post-Keplerian” parameters. Most important among these are: the fractional periastron advance per orbit  $k \equiv \dot{\omega} P_b / 2\pi$ , a dimensionful time-dilation parameter  $\gamma_{\text{timing}}$ , the orbital period derivative  $\dot{P}_b$ , and the “range” and “shape” parameters of the gravitational time delay caused by the companion,  $r$  and  $s$ .

Without assuming any specific theory of gravity, one can phenomenologically analyze the data from any binary pulsar by least-squares fitting the observed sequence of pulse arrival times to the timing formula Eq. (21.18). This fit yields the “measured” values of the parameters  $\{\nu_p, \dot{\nu}_p, \ddot{\nu}_p\}$ ,  $\{p^K\}$ ,  $\{p^{PK}\}$ . Now, each specific relativistic theory of gravity predicts that, for instance,  $k$ ,  $\gamma_{\text{timing}}$ ,  $\dot{P}_b$ ,  $r$  and  $s$  (to quote parameters that have been successfully measured from some binary pulsar data) are some theory-dependent functions of the Keplerian parameters and of the (unknown) masses  $m_1, m_2$  of the pulsar and its companion. For instance, in General Relativity, one finds (with  $M \equiv m_1 + m_2$ ,  $n \equiv 2\pi/P_b$ )

$$\begin{aligned} k^{\text{GR}}(m_1, m_2) &= 3(1 - e^2)^{-1} (G_N M n / c^3)^{2/3} , \\ \gamma_{\text{timing}}^{\text{GR}}(m_1, m_2) &= e n^{-1} (G_N M n / c^3)^{2/3} m_2 (m_1 + 2m_2) / M^2 , \\ \dot{P}_b^{\text{GR}}(m_1, m_2) &= - (192\pi/5) (1 - e^2)^{-7/2} \left( 1 + \frac{73}{24} e^2 + \frac{37}{96} e^4 \right) \\ &\quad \times (G_N M n / c^3)^{5/3} m_1 m_2 / M^2 , \\ r(m_1, m_2) &= G_N m_2 / c^3 , \\ s(m_1, m_2) &= n x (G_N M n / c^3)^{-1/3} M / m_2 . \end{aligned} \quad (21.19)$$

In tensor-scalar theories, each of the functions  $k^{\text{theory}}(m_1, m_2)$ ,  $\gamma_{\text{timing}}^{\text{theory}}(m_1, m_2)$ ,  $\dot{P}_b^{\text{theory}}(m_1, m_2)$ , *etc.*, is modified by quasi-static strong field effects (associated with the self-gravities of the pulsar and its companion), while the particular function  $\dot{P}_b^{\text{theory}}(m_1, m_2)$  is further modified by radiative effects (associated with the spin 0 propagator) [36,45,46].

Let us give some highlights of the current experimental situation. In the first discovered binary pulsar PSR 1913+16 [40,41], it has been

possible to measure with accuracy *three* post-Keplerian parameters:  $k$ ,  $\gamma_{\text{timing}}$  and  $\dot{P}_b$ . The three equations  $k^{\text{measured}} = k^{\text{theory}}(m_1, m_2)$ ,  $\gamma_{\text{timing}}^{\text{measured}} = \gamma_{\text{timing}}^{\text{theory}}(m_1, m_2)$ ,  $\dot{P}_b^{\text{measured}} = \dot{P}_b^{\text{theory}}(m_1, m_2)$  determine, for each given theory, three curves in the two-dimensional mass plane. This yields *one* (combined radiative/strong-field) test of the specified theory, according to whether the three curves meet at one point, as they should. After subtracting a small ( $\sim 10^{-14}$  level in  $\dot{P}_b^{\text{obs}} = (-2.423 \pm 0.001) \times 10^{-12}$ ), but significant, “galactic” perturbing effect (linked to galactic accelerations and to the pulsar proper motion) [47], one finds that General Relativity passes this  $(k - \gamma_{\text{timing}} - \dot{P}_b)_{1913+16}$  test with complete success at the  $10^{-3}$  level [41,48,49]

$$\left[ \frac{\dot{P}_b^{\text{obs}} - \dot{P}_b^{\text{gal}}}{\dot{P}_b^{\text{GR}}[k^{\text{obs}}, \gamma_{\text{timing}}^{\text{obs}}]} \right]_{1913+16} = 0.997 \pm 0.002 . \quad (21.20)$$

Here  $\dot{P}_b^{\text{GR}}[k^{\text{obs}}, \gamma_{\text{timing}}^{\text{obs}}]$  is the result of inserting in  $\dot{P}_b^{\text{GR}}(m_1, m_2)$  the values of the masses predicted by the two equations  $k^{\text{obs}} = k^{\text{GR}}(m_1, m_2)$ ,  $\gamma_{\text{timing}}^{\text{obs}} = \gamma_{\text{timing}}^{\text{GR}}(m_1, m_2)$ . This yields experimental evidence for the reality of gravitational radiation damping forces at the  $(-3 \pm 2) \times 10^{-3}$  level.

The discovery of the binary pulsar PSR 1534+12 [50] has allowed one to measure *five* post-Keplerian parameters:  $k$ ,  $\gamma_{\text{timing}}$ ,  $r$ ,  $s$ , and (with less accuracy)  $\dot{P}_b$  [51,52]. This allows one to obtain *three* (five observables minus two masses) tests of relativistic gravity. Two among these tests probe strong field gravity, without mixing of radiative effects [51]. General Relativity passes all these tests within the measurement accuracy. The most precise of the new, pure strong-field tests is the one obtained by combining the measurements of  $k$ ,  $\gamma_{\text{timing}}$ , and  $s$ . Using the most recent data [52], one finds agreement at the  $(2 \pm 2) \times 10^{-3}$  level:

$$\left[ \frac{s^{\text{obs}}}{s^{\text{GR}}[k^{\text{obs}}, \gamma_{\text{timing}}^{\text{obs}}]} \right]_{1534+12} = 1.002 \pm 0.002 . \quad (21.21)$$

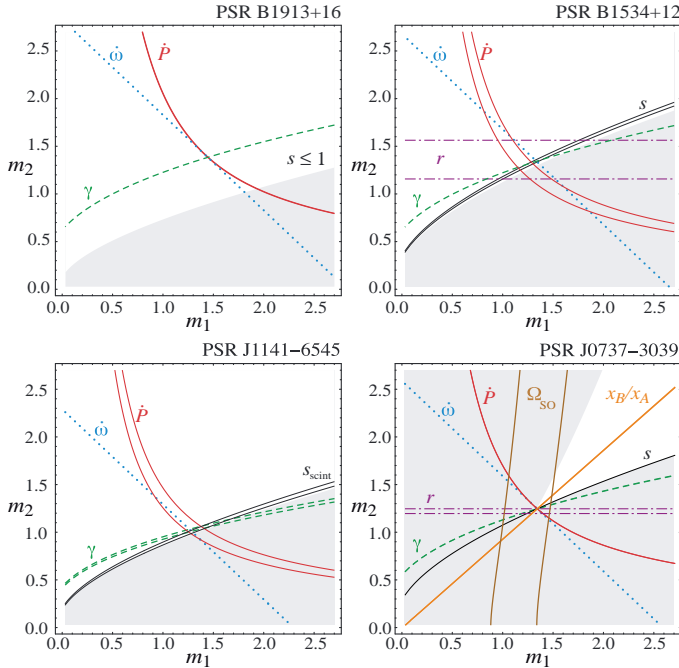
The discovery of the binary pulsar PSR J1141–6545 [53] (whose companion is probably a white dwarf) has allowed one to measure *four* observable parameters:  $k$ ,  $\gamma_{\text{timing}}$ ,  $\dot{P}_b$  [54,55], and the parameter  $s$  [56,55]. The latter parameter (which is equal to the sine of the inclination angle,  $s = \sin i$ ) was consistently measured in two ways: from a scintillation analysis [56], and from timing measurements [55]. General Relativity passes all the corresponding tests within measurement accuracy. See Fig. 21.1 which uses the (more precise) scintillation measurement of  $s = \sin i$ .

The discovery of the remarkable *double* binary pulsar PSR J0737–3039 A and B [57,58] has led to the measurement of *seven* independent parameters [59,60,61]: five of them are the post-Keplerian parameters  $k$ ,  $\gamma_{\text{timing}}$ ,  $r$ ,  $s$  and  $\dot{P}_b$  entering the relativistic timing formula of the fast-spinning pulsar PSR J0737–3039 A, a sixth is the ratio  $R = x_B/x_A$  between the projected semi-major axis of the more slowly spinning companion pulsar PSR J0737–3039 B, and that of PSR J0737–3039 A. [The theoretical prediction for the ratio  $R = x_B/x_A$ , considered as a function of the (inertial) masses  $m_1 = m_A$  and  $m_2 = m_B$ , is  $R^{\text{theory}} = m_1/m_2 + O((v/c)^4)$  [44], independently of the gravitational theory considered.] Finally, the seventh parameter  $\Omega_{\text{SO,B}}$  is the angular rate of (spin-orbit) precession of PSR J0737–3039 B around the total angular momentum [60,61]. These seven measurements give us *five* tests of relativistic gravity [59,62,63]. General Relativity passes all those tests with flying colors (see Fig. 21.1). Let us highlight here two of them (from [63]).

One test is a new confirmation of the reality of gravitational radiation at the  $10^{-3}$  level

$$\left[ \frac{\dot{P}_b^{\text{obs}}}{\dot{P}_b^{\text{GR}}[k^{\text{obs}}, R^{\text{obs}}]} \right]_{0737-3039} = 1.000 \pm 0.001 . \quad (21.22)$$

Another one is a new,  $5 \times 10^{-4}$  level, strong-field confirmation of General Relativity:



**Figure 21.1:** Illustration of the *eleven* tests of relativistic gravity obtained in the four different binary pulsar systems PSR1913+16 (one test), PSR1534+12 (3 tests), PSR J1141–6545 (2 tests), and PSR J0737–3039 A,B (5 tests). Each curve (or strip) in the mass plane corresponds to the interpretation, within General Relativity, of some observable parameter among:  $\dot{P}_b$ ,  $k \equiv \dot{\omega} P_b / 2\pi$ ,  $\gamma_{\text{timing}}$ ,  $s = \sin i$ ,  $\Omega_{\text{SO},B}$  and  $R$ . (Figure updated from [75]; courtesy of G. Esposito-Farèse.)

$$\left[ \frac{s^{\text{obs}}}{s^{\text{GR}}[k^{\text{obs}}, R^{\text{obs}}]} \right]_{0737-3039} = 1.0000 \pm 0.0005. \quad (21.23)$$

Fig. 21.1 illustrates all the tests of strong-field and radiative gravity derived from the above-mentioned binary pulsars:  $(3 - 2 =)$  one test from PSR1913+16,  $(5 - 2 =)$  3 tests from PSR1534+12,  $(4 - 2 =)$  2 tests from PSR J1141–6545, and  $(7 - 2 =)$  5 tests from PSR J0737–3039. [See, also, [64] for additional, less accurate, and partially discrepant, tests of relativistic gravity.]

Data from several nearly circular binary systems (made of a neutron star and a white dwarf) have also led to strong-field confirmations (at the  $4.6 \times 10^{-3}$  level) of the ‘strong equivalence principle,’ *i.e.*, the fact that neutron stars and white dwarfs fall with the same acceleration in the gravitational field of the Galaxy [65,66,67]. The measurements of  $\dot{P}_b$  in some pulsar-white dwarf systems lead to strong constraints on the variation of Newton’s  $G_N$ , and on the existence of gravitational dipole radiation [68,69,70,71,72]. In addition, arrays of millisecond pulsars are sensitive detectors of ultra low frequency gravitational waves ( $f \sim 10^{-9} - 10^{-8}$  Hz) [73]. Such waves might be generated by supermassive black-hole binary systems, by cosmic strings and/or during the inflationary era. The sensitivity of pulsar timing arrays is comparable to predicted gravitational wave signal levels and has recently obtained the most stringent current limit on the energy density of a stochastic relic background of gravitational waves, namely (using standard notation, as in the following *Review* on Big-Bang Cosmology)  $\Omega_{gw}(f)h^2 < 4.2 \times 10^{-10}$  [74].

The constraints on tensor-scalar theories provided by the various binary-pulsar ‘‘experiments’’ have been analyzed in [51,46,75,76,70] and shown to exclude a large portion of the parameter space allowed by solar-system tests. The most stringent tests follow from the measurement of the orbital period decay  $\dot{P}_b$  of the low-eccentricity

8.5-hour pulsar-white dwarf system PSR J1738+0333 with [70]

$$\left[ \dot{P}_b^{\text{obs}} - \dot{P}_b^{\text{gal}} - \dot{P}_b^{\text{GR}} \right]_{1738+0333} = (2.0 \pm 3.7) \times 10^{-15}. \quad (21.24)$$

Asymmetric binary systems are strong emitters of dipolar gravitational radiation in tensor-scalar theories, with  $\dot{P}_b$  scaling (modulo matter-scalar couplings) like  $m_1 m_2 / (m_1 + m_2)^2 (v/c)^3$  ( $\sim 10^{-9}$  for PSR J1738+0333), instead of the smaller quadrupolar radiation  $\dot{P}_b \sim (v/c)^5$  [2,36]. Thereby, the result Eq. (21.24) constrains the basic matter-scalar coupling  $\alpha_0^2$  more strongly, over most of the parameter space, than the best current solar-system limits Eq. (21.16), Eq. (21.17) (namely below the  $10^{-5}$  level) [70]. In the particular case of the Jordan-Fierz-Brans-Dicke theory, the pulsar bound on  $\alpha_0^2$  is (when choosing an equation of state of medium stiffness)  $\alpha_0^2 < 2 \times 10^{-5}$ , which is within a factor two of the Cassini bound Eq. (21.16) (where  $\bar{\gamma} = -2\alpha_0^2 / (1 + \alpha_0^2)$ ).

Measurements over several years of the pulse profiles of various pulsars have detected secular profile changes compatible with the prediction [77] that the general relativistic spin-orbit coupling should cause a secular change in the orientation of the pulsar beam with respect to the line of sight (‘‘geodetic precession’’). Such confirmations of general-relativistic spin-orbit effects were obtained in PSR 1913+16 [78], PSR B1534+12 [52], PSR J1141–6545 [79], PSR J0737–3039 [60,61] and PSR J1906+0746 [80]. In some cases (notably PSR 1913+16 and PSR J1906+0746) the secular change in the orientation of the pulsar beam is expected to lead to the disappearance of the beam (as seen on the Earth) on a human time scale (the second pulsar in the double system PSR J0737–3039 has already disappeared in March 2008 and is expected to reappear around 2035 [61]).

The first observation on September 14, 2015, by the two detectors of the Laser Interferometer Gravitational-wave Observatory (LIGO), of a gravitational-wave signal, and its subsequent analysis by the LIGO/Virgo collaboration [81], has opened up a novel testing ground for relativistic gravity. This transient signal is most readily interpreted as the gravitational-wave signal emitted ( $\sim 400$  Mpc away) by the last  $\sim 3$  inspiralling orbits and the merger of a binary black hole. Thanks to the rather high signal-to-noise ratio ( $\sim 24$ ) of the LIGO observations, one could test consistency with General Relativity in several ways, notably via the good global agreement between the full observed signal and the signal predicted by both analytical [82] and numerical [83] calculations of the gravitational waveform emitted by coalescing black holes. At this early stage, this agreement does not lead to quantitatively accurate tests of relativistic gravity competing with those discussed above. However, this observation has already brought new, deep qualitative confirmations of General Relativity, namely: (i) the first observation, in the wave zone, of gravitational waves; and (ii) the first direct evidence of the existence of black holes via the observation of their merger, followed by an abrupt shut-off of the gravitational-wave signal. Future gravitational-wave observations are expected to be able to probe more deeply relativistic gravity, notably in testing the specific, Einsteinian transverse-traceless quadrupolar nature of gravitational waves, as well as the emission, just after the merger of two black holes, of the characteristic ringing modes of the final, perturbed black hole [84].

The tests considered above have examined the gravitational interaction on scales between a fraction of a millimeter and a few astronomical units. The general relativistic action on light and matter of an external gravitational field have been verified on much larger scales in many gravitational lensing systems [85]. Some tests on cosmological scales are also available [86]. Beyond the various quantitative limits on various parametrized theoretical models discussed in the latter reference, one should remember the massive (strong-field-type) qualitative verification of General Relativity embodied in the fact that relativistic cosmological models give an accurate picture of the Universe over a period during which the spatial metric has been blown up by a gigantic factor, say  $(1+z)^2 \sim 10^{19}$  between Big Bang nucleosynthesis and now.



## 21.4. Conclusions

All present experimental tests are compatible with the predictions of the current “standard” theory of gravitation: Einstein’s General Relativity. The universality of the coupling between matter and gravity (Equivalence Principle) has been verified around the  $10^{-13}$  level. Solar system experiments have tested the weak-field predictions of Einstein’s theory at the few  $10^{-5}$  level. The propagation properties (in the near zone) of relativistic gravity, as well as several of its strong-field aspects, have been verified at the  $10^{-3}$  level (or better) in several binary pulsar experiments. The existence of gravitational waves (in the wave zone), and a direct observational proof of the existence of coalescing black holes, have been obtained by interferometric detectors of gravitational radiation. Recent laboratory experiments have set strong constraints on sub-millimeter modifications of Newtonian gravity. Quantitative confirmations of General Relativity have also been obtained on astrophysical and cosmological scales (though a skeptic might wish to keep in mind the two “dark clouds” of current cosmology, namely the need to assume dark matter and a cosmological constant).

### References:

1. S. Weinberg, *Rev. Mod. Phys.* **61**, 1 (1989).
2. C.M. Will, *Theory and Experiment in Gravitational Physics* (Cambridge University Press, Cambridge, 1993);  
*idem*, *Living Rev. Rel.* **17**, 4 (2014).
3. A.I. Shlyakhter, *Nature* **264**, 340 (1976).
4. T. Damour and F. Dyson, *Nucl. Phys.* **B480**, 37 (1996);  
C.R. Gould, E.I. Sharapov, and S.K. Lamoreaux, *Phys. Rev.* **C74**, 024607 (2006);  
Yu.V. Petrov *et al.*, *Phys. Rev.* **C74**, 064610 (2006).
5. V.V. Flambaum and R.B. Wiringa, *Phys. Rev.* **C79**, 034302 (2009).
6. K.A. Olive *et al.*, *Phys. Rev.* **D66**, 045022 (2002);  
K.A. Olive *et al.*, *Phys. Rev.* **D69**, 027701 (2004).
7. P. Molaro *et al.*, *Astron. & Astrophys.* **555**, A68 (2013).
8. N. Kanekar *et al.*, *Mon. Not. Roy. Astron. Soc.* **448**, no. 1, L104 (2015).
9. J. Bagdonaitė *et al.*, *Phys. Rev. Lett.* **114**, 071301 (2015).
10. S.A. Levshakov *et al.*, *Astron. & Astrophys.* **540**, L9 (2012).
11. T. Rosenband *et al.*, *Science* **319**, 1808 (2008);  
J. Guéna *et al.*, *Phys. Rev. Lett.* **109**, 080801 (2012);  
R. M. Godun *et al.*, *Phys. Rev. Lett.* **113**, 210801 (2014).
12. T.M. Fortier *et al.*, *Phys. Rev. Lett.* **98**, 070801 (2007);  
S. Blatt *et al.*, *Phys. Rev. Lett.* **100**, 140801 (2008);  
T. Dent, *Phys. Rev. Lett.* **101**, 041102 (2008).
13. J.-P. Uzan, *Living Rev. Rel.* **14**, 2 (2011).
14. M. Smiciklas *et al.*, *Phys. Rev. Lett.* **107**, 171604 (2011).
15. S. Liberati, *J. Phys. Conf. Ser.* **631**, no. 1, 012011 (2015).
16. S. Schlamminger *et al.*, *Phys. Rev. Lett.* **100**, 041101 (2008).
17. T.A. Wagner *et al.*, *Class. Quantum Grav.* **29**, 184002 (2012).
18. S. Merlet *et al.*, *Metrologia*, **47**, L9-L11 (2010).
19. D. Schlippert *et al.*, *Phys. Rev. Lett.* **112**, 203002 (2014).
20. J.G. Williams, S.G. Turyshev, and D.H. Boggs, *Class. Quantum Grav.* **29**, 184004 (2012);  
J. Müller, F. Hofmann, and L. Biskupek, *Class. Quantum Grav.* **29**, 184006 (2012).
21. K. Nordtvedt, *Phys. Rev.* **170**, 1186 (1968).
22. E.G. Adelberger *et al.*, *Prog. in Part. Nucl. Phys.* **62**, 102 (2009).
23. R.F.C. Vessot and M.W. Levine, *Gen. Rel. Grav.* **10**, 181 (1978);  
R.F.C. Vessot *et al.*, *Phys. Rev. Lett.* **45**, 2081 (1980).
24. C.W. Chou *et al.*, *Science* **329**, 1630 (2010).
25. I. Ciufolini and E.C. Pavlis, *Nature* **431**, 958 (2004).
26. I. Ciufolini, *Nucl. Phys. Proc. Suppl.* **243-244**, 180 (2013);  
I. Ciufolini *et al.*, *Eur. Phys. J. Plus* **130**, no. 7, 133 (2015).
27. C.W.F. Everitt *et al.*, *Phys. Rev. Lett.* **106**, 221101 (2011).
28. B. Famaey and S. McGaugh, *Living Rev. Rel.* **15**, 10 (2012).
29. C. de Rham, *Living Rev. Rel.* **17**, 7 (2014).
30. E. Berti *et al.*, [arXiv:1501.07274 \[gr-qc\]](https://arxiv.org/abs/1501.07274).
31. T.R. Taylor and G. Veneziano, *Phys. Lett.* **B213**, 450 (1988);  
T. Damour and A.M. Polyakov, *Nucl. Phys.* **B423**, 532 (1994);  
S. Dimopoulos and G. Giudice, *Phys. Lett.* **B379**, 105 (1996);  
I. Antoniadis, S. Dimopoulos, and G. Dvali, *Nucl. Phys.* **B516**, 70 (1998);  
T. Damour and J.F. Donoghue, *Phys. Rev. D* **82**, 084033 (2010).
32. V.A. Rubakov, *Phys. Usp* **44**, 871 (2001);  
R. Maartens and K. Koyama, *Living Rev. Rel.* **13**, 5 (2010).
33. R. D. Newman *et al.*, *Space Sci. Rev.* **148**, 175 (2009).
34. D. J. Kapner *et al.*, *Phys. Rev. Lett.* **98**, 021101 (2007).
35. P. Jordan, *Schwerkraft und Weltall* (Vieweg, Braunschweig, 1955);  
M. Fierz, *Helv. Phys. Acta* **29**, 128 (1956);  
C. Brans and R.H. Dicke, *Phys. Rev.* **124**, 925 (1961).
36. T. Damour and G. Esposito-Farèse, *Class. Quantum Grav.* **9**, 2093 (1992).
37. J. Khoury and A. Weltman, *Phys. Rev. Lett.* **93**, 171104 (2004).
38. B. Bertotti, L. Iess, and P. Tortora, *Nature* **425**, 374 (2003).
39. A. Fienga *et al.*, *Cel. Mech. Dyn. Astr.* **123**, Issue 2, 1 (2015).
40. R.A. Hulse, *Rev. Mod. Phys.* **66**, 699 (1994).
41. J.H. Taylor, *Rev. Mod. Phys.* **66**, 711 (1994).
42. G.D. Moore and A.E. Nelson, *JHEP* **0109**, 023 (2001).
43. T. Damour and N. Deruelle, *Phys. Lett.* **A87**, 81 (1981);  
T. Damour, *C.R. Acad. Sci. Paris* **294**, 1335 (1982).
44. T. Damour and N. Deruelle, *Ann. Inst. H. Poincaré A*, **44**, 263 (1986);  
Damour and J.H. Taylor, *Phys. Rev.* **D45**, 1840 (1992).
45. C.M. Will and H.W. Zaglauer, *Astrophys. J.* **346**, 366 (1989).
46. T. Damour and G. Esposito-Farèse, *Phys. Rev.* **D54**, 1474 (1996);  
*idem*, *Phys. Rev.* **D58**, 042001 (1998).
47. T. Damour and J.H. Taylor, *Astrophys. J.* **366**, 501 (1991).
48. J.M. Weisberg and J.H. Taylor, in *Radio Pulsars, ASP Conference Series* **328**, 25 (2005); [astro-ph/0407149](https://arxiv.org/abs/astro-ph/0407149).
49. J.M. Weisberg, D.J. Nice, and J.H. Taylor, *Astrophys. J.* **722**, 1030 (2010).
50. A. Wolszczan, *Nature* **350**, 688 (1991).
51. J.H. Taylor *et al.*, *Nature* **355**, 132 (1992).
52. E. Fonseca, I.H. Stairs and S.E. Thorsett, *Astrophys. J.* **787**, 82 (2014).
53. V.M. Kaspi *et al.*, *Astrophys. J.* **528**, 445 (2000).
54. M. Bailes *et al.*, *Astrophys. J.* **595**, L49 (2003).
55. J.P.W. Verbiest, N.D.R. Bhat, and M. Bailes, ISBN 978-981-4374-51-4. Singapore: World Scientific (2012) 1571 [[arXiv:1210.0224](https://arxiv.org/abs/1210.0224) [[astro-ph](https://arxiv.org/abs/astro-ph).HE]].
56. S.M. Ord *et al.*, *Astrophys. J.* **574**, L75 (2002).
57. M. Burgay *et al.*, *Nature* **426**, 531 (2003).
58. A.G. Lyne *et al.*, *Science* **303**, 1153 (2004).
59. M. Kramer *et al.*, *Science* **314**, 97 (2006).
60. R.P. Breton *et al.*, *Science* **321**, 104 (2008).
61. B. Perera *et al.*, *AIP Conf. Proc.* **1357**, 105 (2011) [*Astrophys. J.* **721**, 1193 (2010)].
62. M. Kramer and N. Wex, *Class. Quantum Grav.* **26**, 073001 (2009).
63. M. Kramer, in *Neutron Stars and Pulsars: Challenges and Opportunities after 80 Years; Proceedings of the International Astronomical Union Symposium S291, 2012*, J. van Leeuwen, ed. (Cambridge University Press, 2013), pp 19-26, [[arXiv:1211.2457](https://arxiv.org/abs/1211.2457) [[astro-ph](https://arxiv.org/abs/astro-ph).HE]].
64. R.D. Ferdman *et al.*, *Mon. Not. Roy. Astron. Soc.* **443**, no. 3, 2183 (2014).
65. T. Damour and G. Schäfer, *Phys. Rev. Lett.* **66**, 2549 (1991).
66. M.E. Gonzalez *et al.*, *Astrophys. J.* **743**, 102 (2011).
67. P.C.C. Freire, M. Kramer, and N. Wex, *Class. Quantum Grav.* **29**, 184007 (2012).
68. J.P.W. Verbiest *et al.*, *Astrophys. J.* **679**, 675 (2008).
69. K. Lazaridis *et al.*, *Mon. Not. Roy. Astron. Soc.* **400**, 805 (2009).
70. P.C.C. Freire *et al.*, *Mon. Not. Roy. Astron. Soc.* **423**, 3328 (2012).
71. J. Antoniadis *et al.*, *Science* **340**, 6131 (2013).
72. W.W. Zhu *et al.*, *Astrophys. J.* **809**, 41 (2015).
73. R.S. Foster and D.C. Backer *Astrophys. J.* **361**, 300 (1990).
74. L. Lentati *et al.*, *Mon. Not. Roy. Astron. Soc.* **453**, 2576 (2015);  
Z. Arzoumanian *et al.* [NANOGrav Collaboration], [arXiv:1508.03024](https://arxiv.org/abs/1508.03024) [[astro-ph](https://arxiv.org/abs/astro-ph).GA].

75. G. Esposito-Farèse, in *Proceedings of the 10th Marcel Grossmann Meeting on Recent Developments in Theoretical and Experimental General Relativity*, edited by M. Novello *et al.*, (World Scientific, 2006), part A, pp 647-666.
76. G. Esposito-Farèse, in *Mass and Motion in General Relativity*, eds L. Blanchet *et al.*, series Fundam. Theor. Phys. **162** (Springer, Dordrecht, 2011) 461-489.
77. T. Damour and R. Ruffini, C. R. Acad. Sc. Paris **279**, série A, 971 (1974);  
B.M. Barker and R.F. O'Connell, Phys. Rev. **D12**, 329 (1975).
78. M. Kramer, Astrophys. J. **509**, 856 (1998);  
J.M. Weisberg and J.H. Taylor, Astrophys. J. **576**, 942 (2002).
79. R.N. Manchester *et al.*, Astrophys. J. **710**, 1694 (2010).
80. J. van Leeuwen *et al.*, Astrophys. J. **798**, 118 (2015).
81. B.P. Abbott *et al.*, Phys. Rev. Lett. **116**, 061102 (2016).
82. A. Buonanno and T. Damour, Phys. Rev. **D62**, 064015 (2000).
83. F. Pretorius, Phys. Rev. Lett. **95**, 121101 (2005);  
M. Campanelli *et al.*, Phys. Rev. Lett. **96**, 111101 (2006);  
J.G. Baker *et al.*, Phys. Rev. Lett. **96**, 111102 (2006).
84. C.V. Vishveshwara, Nature **227**, 936 (1970).
85. A.S. Bolton, S. Rappaport, and S. Burles, Phys. Rev. **D74**, 061501 (2006);  
J. Schwab, A.S. Bolton, and S.A. Rappaport, Astrophys. J. **708**, 750 (2010).
86. J.-P. Uzan, Gen. Rel. Grav. **42**, 2219 (2010);  
R. Reyes *et al.*, Nature **464**, 256 (2010);  
B. Jain and J. Khoury, Ann. Phys. **325**, 1479 (2010);  
S.F. Daniel, *et al.*, Phys. Rev. **D81**, 123508 (2010).

## 22. BIG-BANG COSMOLOGY

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### 22.1. Introduction to Standard Big-Bang Model

The observed expansion of the Universe [1–3] is a natural (almost inevitable) result of any homogeneous and isotropic cosmological model based on general relativity. However, by itself, the Hubble expansion does not provide sufficient evidence for what we generally refer to as the Big-Bang model of cosmology. While general relativity is in principle capable of describing the cosmology of any given distribution of matter, it is extremely fortunate that our Universe appears to be homogeneous and isotropic on large scales. Together, homogeneity and isotropy allow us to extend the Copernican Principle to the Cosmological Principle, stating that all spatial positions in the Universe are essentially equivalent.

The formulation of the Big-Bang model began in the 1940s with the work of George Gamow and his collaborators, Alpher and Herman. In order to account for the possibility that the abundances of the elements had a cosmological origin, they proposed that the early Universe which was once very hot and dense (enough so as to allow for the nucleosynthetic processing of hydrogen), and has expanded and cooled to its present state [4,5]. In 1948, Alpher and Herman predicted that a direct consequence of this model is the presence of a relic background radiation with a temperature of order a few K [6,7]. Of course this radiation was observed 16 years later as the microwave background radiation [8]. Indeed, it was the observation of the 3 K background radiation that singled out the Big-Bang model as the prime candidate to describe our Universe. Subsequent work on Big-Bang nucleosynthesis further confirmed the necessity of our hot and dense past. (See Sec. 22.3.7 for a brief discussion of BBN and the review on BBN—Sec. 24 of this *Review* for a detailed discussion of BBN.) These relativistic cosmological models face severe problems with their initial conditions, to which the best modern solution is inflationary cosmology, discussed in Sec. 22.3.5. (See the upcoming review on inflation of this *Review* for a detailed discussion of inflation.) If correct, these ideas would strictly render the term ‘Big Bang’ redundant, since it was first coined by Hoyle to represent a criticism of the lack of understanding of the initial conditions.

#### 22.1.1. The Robertson-Walker Universe :

The observed homogeneity and isotropy enable us to describe the overall geometry and evolution of the Universe in terms of two cosmological parameters accounting for the spatial curvature and the overall expansion (or contraction) of the Universe. These two quantities appear in the most general expression for a space-time metric which has a (3D) maximally symmetric subspace of a 4D space-time, known as the Robertson-Walker metric:

$$ds^2 = dt^2 - R^2(t) \left[ \frac{dr^2}{1 - kr^2} + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \right]. \quad (22.1)$$

Note that we adopt  $c = 1$  throughout. By rescaling the radial coordinate  $r$  we can choose the curvature constant  $k$  to take only the discrete values  $+1$ ,  $-1$ , or  $0$  corresponding to closed, open, or spatially flat geometries. In this case, it is often more convenient to re-express the metric as

$$ds^2 = dt^2 - R^2(t) \left[ d\chi^2 + S_k^2(\chi) (d\theta^2 + \sin^2 \theta d\phi^2) \right], \quad (22.2)$$

where the function  $S_k(\chi)$  is  $(\sin \chi, \chi, \sinh \chi)$  for  $k = (+1, 0, -1)$ . The coordinate  $r$  [in Eq. (22.1)] and the ‘angle’  $\chi$  (in Eq. (22.2)) are both dimensionless; the dimensions are carried by  $R(t)$ , which is the cosmological scale factor which determines proper distances in terms of the comoving coordinates. A common alternative is to define a dimensionless scale factor,  $a(t) = R(t)/R_0$ , where  $R_0 \equiv R(t_0)$  is  $R$  at the present epoch. It is also sometimes convenient to define a dimensionless or conformal time coordinate,  $\eta$ , by  $d\eta = dt/R(t)$ . Along constant spatial sections, the proper time is defined by the time coordinate,  $t$ . Similarly, for  $dt = d\theta = d\phi = 0$ , the proper distance is given by  $R(t)\chi$ . For standard texts on cosmological models see *e.g.*, Refs. [9–16].

#### 22.1.2. The redshift :

The cosmological redshift is a direct consequence of the Hubble expansion, determined by  $R(t)$ . A local observer detecting light from a distant emitter sees a redshift in frequency. We can define the redshift as

$$z \equiv \frac{\nu_1 - \nu_2}{\nu_2} \simeq v_{12}, \quad (22.3)$$

where  $\nu_1$  is the frequency of the emitted light,  $\nu_2$  is the observed frequency and  $v_{12}$  is the relative velocity between the emitter and the observer. While the definition,  $z = (\nu_1 - \nu_2)/\nu_2$  is valid on all distance scales, relating the redshift to the relative velocity in this simple way is only true on small scales (*i.e.*, less than cosmological scales) such that the expansion velocity is non-relativistic. For light signals, we can use the metric given by Eq. (22.1) and  $ds^2 = 0$  to write

$$v_{12} = \dot{R} \delta r = \frac{\dot{R}}{R} \delta t = \frac{\delta R}{R} = \frac{R_2 - R_1}{R_1}, \quad (22.4)$$

where  $\delta r(\delta t)$  is the radial coordinate (temporal) separation between the emitter and observer. Noting that physical distance,  $D$ , is  $R\delta r$  or  $\delta t$ , Eq. (22.4) gives us Hubble’s law,  $v = HD$ . In addition, we obtain the simple relation between the redshift and the scale factor

$$1 + z = \frac{\nu_1}{\nu_2} = \frac{R_2}{R_1}. \quad (22.5)$$

This result does not depend on the non-relativistic approximation.

#### 22.1.3. The Friedmann equations of motion :

The cosmological equations of motion are derived from Einstein’s equations

$$\mathcal{R}_{\mu\nu} - \frac{1}{2}g_{\mu\nu}\mathcal{R} = 8\pi G_N T_{\mu\nu} + \Lambda g_{\mu\nu}. \quad (22.6)$$

Gliner [17] and Zeldovich [18] have pioneered the modern view, in which the  $\Lambda$  term is set on the rhs and interpreted as an effective energy-momentum tensor  $T_{\mu\nu}$  for the vacuum of  $\Lambda g_{\mu\nu}/8\pi G_N$ . It is common to assume that the matter content of the Universe is a perfect fluid, for which

$$T_{\mu\nu} = -pg_{\mu\nu} + (p + \rho)u_\mu u_\nu, \quad (22.7)$$

where  $g_{\mu\nu}$  is the space-time metric described by Eq. (22.1),  $p$  is the isotropic pressure,  $\rho$  is the energy density and  $u = (1, 0, 0, 0)$  is the velocity vector for the isotropic fluid in co-moving coordinates. With the perfect fluid source, Einstein’s equations lead to the Friedmann equations

$$H^2 \equiv \left( \frac{\dot{R}}{R} \right)^2 = \frac{8\pi G_N \rho}{3} - \frac{k}{R^2} + \frac{\Lambda}{3}, \quad (22.8)$$

and

$$\frac{\ddot{R}}{R} = \frac{\Lambda}{3} - \frac{4\pi G_N}{3}(\rho + 3p), \quad (22.9)$$

where  $H(t)$  is the Hubble parameter and  $\Lambda$  is the cosmological constant. The first of these is sometimes called the Friedmann equation. Energy conservation via  $T^{\mu\nu}_{;\mu} = 0$ , leads to a third useful equation [which can also be derived from Eq. (22.8) and Eq. (22.9)]

$$\dot{\rho} = -3H(\rho + p). \quad (22.10)$$

Eq. (22.10) can also be simply derived as a consequence of the first law of thermodynamics.

Eq. (22.8) has a simple classical mechanical analog if we neglect (for the moment) the cosmological term  $\Lambda$ . By interpreting  $-k/R^2$  Newtonianly as a ‘total energy’, then we see that the evolution of the Universe is governed by a competition between the potential energy,  $8\pi G_N \rho/3$ , and the kinetic term  $(\dot{R}/R)^2$ . For  $\Lambda = 0$ , it is clear that the Universe must be expanding or contracting (except at the turning point prior to collapse in a closed Universe). The ultimate fate of the Universe is determined by the curvature constant  $k$ . For  $k = +1$ , the Universe will recollapse in a finite time, whereas for  $k = 0, -1$ , the Universe will expand indefinitely. These simple conclusions can be altered when  $\Lambda \neq 0$  or more generally with some component with  $(\rho + 3p) < 0$ .

#### 22.1.4. Definition of cosmological parameters :

In addition to the Hubble parameter, it is useful to define several other measurable cosmological parameters. The Friedmann equation can be used to define a critical density such that  $k = 0$  when  $\Lambda = 0$ ,

$$\begin{aligned}\rho_c &\equiv \frac{3H^2}{8\pi G_N} = 1.88 \times 10^{-26} h^2 \text{ kg m}^{-3} \\ &= 1.05 \times 10^{-5} h^2 \text{ GeV cm}^{-3},\end{aligned}\quad (22.11)$$

where the scaled Hubble parameter,  $h$ , is defined by

$$\begin{aligned}H &\equiv 100 h \text{ km s}^{-1} \text{ Mpc}^{-1} \\ \Rightarrow H^{-1} &= 9.78 h^{-1} \text{ Gyr} \\ &= 2998 h^{-1} \text{ Mpc}.\end{aligned}\quad (22.12)$$

The cosmological density parameter  $\Omega_{\text{tot}}$  is defined as the energy density relative to the critical density,

$$\Omega_{\text{tot}} = \rho/\rho_c. \quad (22.13)$$

Note that one can now rewrite the Friedmann equation as

$$k/R^2 = H^2(\Omega_{\text{tot}} - 1). \quad (22.14)$$

From Eq. (22.14), one can see that when  $\Omega_{\text{tot}} > 1$ ,  $k = +1$  and the Universe is closed, when  $\Omega_{\text{tot}} < 1$ ,  $k = -1$  and the Universe is open, and when  $\Omega_{\text{tot}} = 1$ ,  $k = 0$ , and the Universe is spatially flat.

It is often necessary to distinguish different contributions to the density. It is therefore convenient to define present-day density parameters for pressureless matter ( $\Omega_m$ ) and relativistic particles ( $\Omega_r$ ), plus the quantity  $\Omega_\Lambda = \Lambda/3H^2$ . In more general models, we may wish to drop the assumption that the vacuum energy density is constant, and we therefore denote the present-day density parameter of the vacuum by  $\Omega_v$ . The Friedmann equation then becomes

$$k/R_0^2 = H_0^2(\Omega_m + \Omega_r + \Omega_v - 1), \quad (22.15)$$

where the subscript 0 indicates present-day values. Thus, it is the sum of the densities in matter, relativistic particles, and vacuum that determines the overall sign of the curvature. Note that the quantity  $-k/R_0^2 H_0^2$  is sometimes referred to as  $\Omega_k$ . This usage is unfortunate: it encourages one to think of curvature as a contribution to the energy density of the Universe, which is not correct.

#### 22.1.5. Standard Model solutions :

Much of the history of the Universe in the standard Big-Bang model can be easily described by assuming that either matter or radiation dominates the total energy density. During inflation and again today the expansion rate for the Universe is accelerating, and domination by a cosmological constant or some other form of dark energy should be considered. In the following, we shall delineate the solutions to the Friedmann equation when a single component dominates the energy density. Each component is distinguished by an equation of state parameter  $w = p/\rho$ .

##### 22.1.5.1. Solutions for a general equation of state:

Let us first assume a general equation of state parameter for a single component,  $w$  which is constant. In this case, Eq. (22.10) can be written as  $\dot{\rho} = -3(1+w)\rho\dot{R}/R$  and is easily integrated to yield

$$\rho \propto R^{-3(1+w)}. \quad (22.16)$$

Note that at early times when  $R$  is small, the less singular curvature term  $k/R^2$  in the Friedmann equation can be neglected so long as  $w > -1/3$ . Curvature domination occurs at rather late times (if a cosmological constant term does not dominate sooner). For  $w \neq -1$ , one can insert this result into the Friedmann equation Eq. (22.8), and if one neglects the curvature and cosmological constant terms, it is easy to integrate the equation to obtain,

$$R(t) \propto t^{2/[3(1+w)]}. \quad (22.17)$$

##### 22.1.5.2. A Radiation-dominated Universe:

In the early hot and dense Universe, it is appropriate to assume an equation of state corresponding to a gas of radiation (or relativistic particles) for which  $w = 1/3$ . In this case, Eq. (22.16) becomes  $\rho \propto R^{-4}$ . The ‘extra’ factor of  $1/R$  is due to the cosmological redshift; not only is the number density of particles in the radiation background decreasing as  $R^{-3}$  since volume scales as  $R^3$ , but in addition, each particle’s energy is decreasing as  $E \propto \nu \propto R^{-1}$ . Similarly, one can substitute  $w = 1/3$  into Eq. (22.17) to obtain

$$R(t) \propto t^{1/2}; \quad H = 1/2t. \quad (22.18)$$

##### 22.1.5.3. A Matter-dominated Universe:

At relatively late times, non-relativistic matter eventually dominates the energy density over radiation (see Sec. 22.3.8). A pressureless gas ( $w = 0$ ) leads to the expected dependence  $\rho \propto R^{-3}$  from Eq. (22.16) and, if  $k = 0$ , we get

$$R(t) \propto t^{2/3}; \quad H = 2/3t. \quad (22.19)$$

##### 22.1.5.4. A Universe dominated by vacuum energy:

If there is a dominant source of vacuum energy,  $V_0$ , it would act as a cosmological constant with  $\Lambda = 8\pi G_N V_0$  and equation of state  $w = -1$ . In this case, the solution to the Friedmann equation is particularly simple and leads to an exponential expansion of the Universe:

$$R(t) \propto e^{\sqrt{\Lambda/3}t}. \quad (22.20)$$

A key parameter is the equation of state of the vacuum,  $w \equiv p/\rho$ : this need not be the  $w = -1$  of  $\Lambda$ , and may not even be constant [19–21]. There is now much interest in the more general possibility of a dynamically evolving vacuum energy, for which the name ‘dark energy’ has become commonly used. A variety of techniques exist whereby the vacuum density as a function of time may be measured, usually expressed as the value of  $w$  as a function of epoch [22,23]. The best current measurement for the equation of state (assumed constant, but without assuming zero curvature) is  $w = -0.97 \pm 0.05$  [24]. Unless stated otherwise, we will assume that the vacuum energy is a cosmological constant with  $w = -1$  exactly.

The presence of vacuum energy can dramatically alter the fate of the Universe. For example, if  $\Lambda < 0$ , the Universe will eventually recollapse independent of the sign of  $k$ . For large values of  $\Lambda > 0$  (larger than the Einstein static value needed to halt any cosmological expansion or contraction), even a closed Universe will expand forever. One way to quantify this is the deceleration parameter,  $q_0$ , defined as

$$q_0 = - \left. \frac{R\ddot{R}}{\dot{R}^2} \right|_0 = \frac{1}{2}\Omega_m + \Omega_r + \frac{(1+3w)}{2}\Omega_v. \quad (22.21)$$

This equation shows us that  $w < -1/3$  for the vacuum may lead to an accelerating expansion. To the continuing astonishment of cosmologists, such an effect has been observed; one piece of direct evidence is the Supernova Hubble diagram [25–30] (see Fig. 22.1 below); current data indicate that vacuum energy is indeed the largest contributor to the cosmological density budget, with  $\Omega_v = 0.692 \pm 0.012$  and  $\Omega_m = 0.308 \pm 0.012$  if  $k = 0$  is assumed (Planck) [31].

The existence of this constituent is without doubt the greatest puzzle raised by the current cosmological model; the final section of this review discusses some of the ways in which the vacuum-energy problem is being addressed. For more details, see the review on Dark Energy—Sec. 27

## 22.2. Introduction to Observational Cosmology

### 22.2.1. Fluxes, luminosities, and distances :

The key quantities for observational cosmology can be deduced quite directly from the metric.

(1) The *proper* transverse size of an object seen by us to subtend an angle  $d\psi$  is its comoving size  $d\psi S_k(\chi)$  times the scale factor at the time of emission:

$$d\ell = d\psi R_0 S_k(\chi)/(1+z) . \quad (22.22)$$

(2) The apparent flux density of an object is deduced by allowing its photons to flow through a sphere of current radius  $R_0 S_k(\chi)$ ; but photon energies and arrival rates are redshifted, and the bandwidth  $d\nu$  is reduced. The observed photons at frequency  $\nu_0$  were emitted at frequency  $\nu_0(1+z)$ , so the flux density is the luminosity at this frequency, divided by the total area, divided by  $1+z$ :

$$S_\nu(\nu_0) = \frac{L_\nu([1+z]\nu_0)}{4\pi R_0^2 S_k^2(\chi)(1+z)} . \quad (22.23)$$

These relations lead to the following common definitions:

$$\begin{aligned} \text{angular-diameter distance: } D_A &= (1+z)^{-1} R_0 S_k(\chi) \\ \text{luminosity distance: } D_L &= (1+z) R_0 S_k(\chi) . \end{aligned} \quad (22.24)$$

These distance-redshift relations are expressed in terms of observables by using the equation of a null radial geodesic ( $R(t)d\chi = dt$ ) plus the Friedmann equation:

$$\begin{aligned} R_0 d\chi &= \frac{1}{H(z)} dz = \frac{1}{H_0} \left[ (1-\Omega_m - \Omega_v - \Omega_r)(1+z)^2 \right. \\ &\quad \left. + \Omega_v(1+z)^{3+3w} + \Omega_m(1+z)^3 + \Omega_r(1+z)^4 \right]^{-1/2} dz . \end{aligned} \quad (22.25)$$

The main scale for the distance here is the Hubble length,  $1/H_0$ .

The flux density is the product of the specific intensity  $I_\nu$  and the solid angle  $d\Omega$  subtended by the source:  $S_\nu = I_\nu d\Omega$ . Combining the angular size and flux-density relations thus gives the relativistic version of surface-brightness conservation:

$$I_\nu(\nu_0) = \frac{B_\nu([1+z]\nu_0)}{(1+z)^3} , \quad (22.26)$$

where  $B_\nu$  is surface brightness (luminosity emitted into unit solid angle per unit area of source). We can integrate over  $\nu_0$  to obtain the corresponding total or bolometric formula:

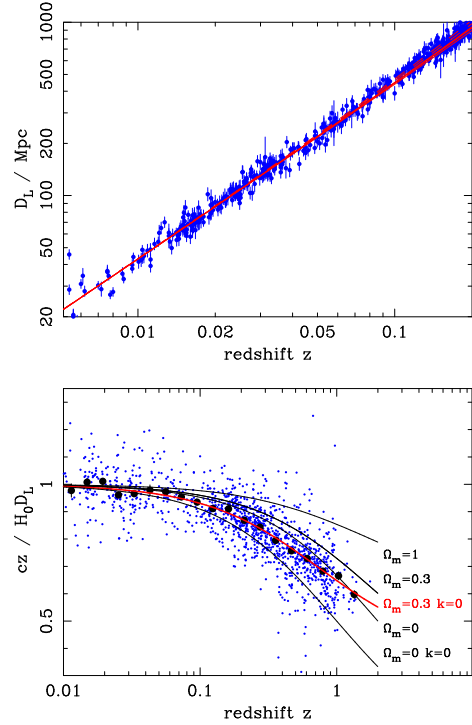
$$I_{\text{tot}} = \frac{B_{\text{tot}}}{(1+z)^4} . \quad (22.27)$$

This cosmology-independent form expresses Liouville's Theorem: photon phase-space density is conserved along rays.

### 22.2.2. Distance data and geometrical tests of cosmology :

In order to confront these theoretical predictions with data, we have to bridge the divide between two extremes. Nearby objects may have their distances measured quite easily, but their radial velocities are dominated by deviations from the ideal Hubble flow, which typically have a magnitude of several hundred  $\text{km s}^{-1}$ . On the other hand, objects at redshifts  $z \gtrsim 0.01$  will have observed recessional velocities that differ from their ideal values by  $\lesssim 10\%$ , but absolute distances are much harder to supply in this case. The traditional solution to this problem is the construction of the distance ladder: an interlocking set of methods for obtaining relative distances between various classes of object, which begins with absolute distances at the 10 to 100 pc level, and terminates with galaxies at significant redshifts. This is reviewed in the review on Cosmological Parameters—Sec. 25 of this *Review*.

By far the most exciting development in this area has been the use of type Ia Supernovae (SNe), which now allow measurement of relative distances with 5% precision. In combination with Cepheid data from the HST and a direct geometrical distance to the maser galaxy



**Figure 22.1:** The type Ia supernova Hubble diagram, based on over 1200 publicly available supernova distance estimates [28–30]. The first panel shows that for  $z \ll 1$  the large-scale Hubble flow is indeed linear and uniform; the second panel shows an expanded scale, with the linear trend divided out, and with the redshift range extended to show how the Hubble law becomes nonlinear. ( $\Omega_r = 0$  is assumed.) Larger points with errors show median values in redshift bins. Comparison with the prediction of Friedmann models appears to favor a vacuum-dominated Universe.

NGC4258, SNe results extend the distance ladder to the point where deviations from uniform expansion are negligible, leading to the best existing direct value for  $H_0$ :  $72.0 \pm 3.0 \text{ km s}^{-1} \text{ Mpc}^{-1}$  [32]. Better still, the analysis of high- $z$  SNe has allowed a simple and direct test of cosmological geometry to be carried out: as shown in Fig. 22.1 and Fig. 22.2, supernova data and measurements of microwave-background anisotropies strongly favor a  $k = 0$  model dominated by vacuum energy. (See the review on Cosmological Parameters—Sec. 25 of this *Review* for a more comprehensive review of Hubble parameter determinations.)

### 22.2.3. Age of the Universe :

The most striking conclusion of relativistic cosmology is that the Universe has not existed forever. The dynamical result for the age of the Universe may be written as

$$\begin{aligned} H_0 t_0 &= \int_0^\infty \frac{dz}{(1+z)H(z)} \\ &= \int_0^\infty \frac{dz}{(1+z) [(1+z)^2(1+\Omega_m z) - z(2+z)\Omega_v]^{1/2}} , \end{aligned} \quad (22.28)$$

where we have neglected  $\Omega_r$  and chosen  $w = -1$ . Over the range of interest ( $0.1 \lesssim \Omega_m \lesssim 1$ ,  $|\Omega_v| \lesssim 1$ ), this exact answer may be approximated to a few % accuracy by

$$H_0 t_0 \simeq \frac{2}{3} (0.7\Omega_m + 0.3 - 0.3\Omega_v)^{-0.3} . \quad (22.29)$$

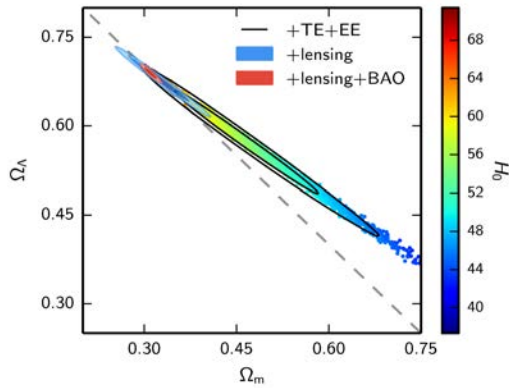
For the special case that  $\Omega_m + \Omega_v = 1$ , the integral in Eq. (22.28) can be expressed analytically as

$$H_0 t_0 = \frac{2}{3\sqrt{\Omega_v}} \ln \frac{1+\sqrt{\Omega_v}}{\sqrt{1-\Omega_v}} \quad (\Omega_m < 1) . \quad (22.30)$$

The most accurate means of obtaining ages for astronomical objects is based on the natural clocks provided by radioactive decay. The use of these clocks is complicated by a lack of knowledge of the initial conditions of the decay. In the Solar System, chemical fractionation of different elements helps pin down a precise age for the pre-Solar nebula of 4.6 Gyr, but for stars it is necessary to attempt an a priori calculation of the relative abundances of nuclei that result from supernova explosions. In this way, a lower limit for the age of stars in the local part of the Milky Way of about 11 Gyr is obtained [34,35].

The other major means of obtaining cosmological age estimates is based on the theory of stellar evolution. In principle, the main-sequence turnoff point in the color-magnitude diagram of a globular cluster should yield a reliable age. However, these have been controversial owing to theoretical uncertainties in the evolution model, as well as observational uncertainties in the distance, dust extinction, and metallicity of clusters. The present consensus favors ages for the oldest clusters of about 12 Gyr [36,37].

These methods are all consistent with the age deduced from studies of structure formation, using the microwave background and large-scale structure:  $t_0 = 13.80 \pm 0.04$  Gyr [31], where the extra accuracy comes at the price of assuming the Cold Dark Matter model to be true.



**Figure 22.2:** Likelihood-based probability densities on the plane  $\Omega_\Lambda$  (i.e.,  $\Omega_v$  assuming  $w = -1$ ) vs  $\Omega_m$ . The colored locus derives from Planck [31] and shows that the CMB alone requires a flat universe  $\Omega_v + \Omega_m \simeq 1$  if the Hubble constant is not too high. The SNe Ia results [33] very nearly constrain the orthogonal combination  $\Omega_v - \Omega_m$ , and the intersection of these constraints directly favors a flat model with  $\Omega_m \simeq 0.3$ , as does the measurement of the Baryon Acoustic Oscillation lengthscale (for which a joint constraint is shown on this plot). The CMB alone is capable of breaking the degeneracy with  $H_0$  by using the measurements of gravitational lensing that can be made with modern high-resolution CMB data.

#### 22.2.4. Horizon, isotropy, flatness problems :

For photons, the radial equation of motion is just  $c dt = R d\chi$ . How far can a photon get in a given time? The answer is clearly

$$\Delta\chi = \int_{t_1}^{t_2} \frac{dt}{R(t)} \equiv \Delta\eta, \quad (22.31)$$

i.e., just the interval of conformal time. We can replace  $dt$  by  $dR/\dot{R}$ , which the Friedmann equation says is  $\propto dR/\sqrt{\rho R^2}$  at early times. Thus, this integral converges if  $\rho R^2 \rightarrow \infty$  as  $t_1 \rightarrow 0$ , otherwise it diverges. Provided the equation of state is such that  $\rho$  changes faster than  $R^{-2}$ , light signals can only propagate a finite distance between the Big Bang and the present; there is then said to be a particle horizon. Such a horizon therefore exists in conventional Big-Bang models, which are dominated by radiation ( $\rho \propto R^{-4}$ ) at early times.

At late times, the integral for the horizon is largely determined by

the matter-dominated phase, for which

$$D_H = R_0 \chi_H \equiv R_0 \int_0^{t(z)} \frac{dt}{R(t)} \simeq \frac{6000}{\sqrt{\Omega_m z}} h^{-1} \text{Mpc} \quad (z \gg 1). \quad (22.32)$$

The horizon at the time of formation of the microwave background ('last scattering':  $z \simeq 1100$ ) was thus of order 100 Mpc in size, subtending an angle of about  $1^\circ$ . Why then are the large number of causally disconnected regions we see on the microwave sky all at the same temperature? The Universe is very nearly isotropic and homogeneous, even though the initial conditions appear not to permit such a state to be constructed.

A related problem is that the  $\Omega = 1$  Universe is unstable:

$$\Omega(a) - 1 = \frac{\Omega - 1}{1 - \Omega + \Omega_v a^2 + \Omega_m a^{-1} + \Omega_r a^{-2}}, \quad (22.33)$$

where  $\Omega$  with no subscript is the total density parameter, and  $a(t) = R(t)/R_0$ . This requires  $\Omega(t)$  to be unity to arbitrary precision as the initial time tends to zero; a universe of non-zero curvature today requires very finely tuned initial conditions.

### 22.3. The Hot Thermal Universe

#### 22.3.1. Thermodynamics of the early Universe :

As alluded to above, we expect that much of the early Universe can be described by a radiation-dominated equation of state. In addition, through much of the radiation-dominated period, thermal equilibrium is established by the rapid rate of particle interactions relative to the expansion rate of the Universe (see Sec. 22.3.3 below). In equilibrium, it is straightforward to compute the thermodynamic quantities,  $\rho$ ,  $p$ , and the entropy density,  $s$ . In general, the energy density for a given particle type  $i$  can be written as

$$\rho_i = \int E_i dn_{q_i}, \quad (22.34)$$

with the density of states given by

$$dn_{q_i} = \frac{g_i}{2\pi^2} (\exp[(E_{q_i} - \mu_i)/T_i] \pm 1)^{-1} q_i^2 dq_i, \quad (22.35)$$

where  $g_i$  counts the number of degrees of freedom for particle type  $i$ ,  $E_{q_i}^2 = m_i^2 + q_i^2$ ,  $\mu_i$  is the chemical potential, and the  $\pm$  corresponds to either Fermi or Bose statistics. Similarly, we can define the pressure of a perfect gas as

$$p_i = \frac{1}{3} \int \frac{q_i^2}{E_i} dn_{q_i}. \quad (22.36)$$

The number density of species  $i$  is simply

$$n_i = \int dn_{q_i}, \quad (22.37)$$

and the entropy density is

$$s_i = \frac{\rho_i + p_i - \mu_i n_i}{T_i}. \quad (22.38)$$

In the Standard Model, a chemical potential is often associated with baryon number, and since the net baryon density relative to the photon density is known to be very small (of order  $10^{-10}$ ), we can neglect any such chemical potential when computing total thermodynamic quantities.

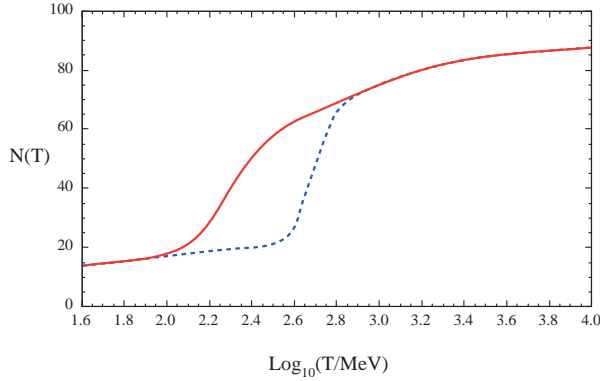
For photons, we can compute all of the thermodynamic quantities rather easily. Taking  $g_i = 2$  for the 2 photon polarization states, we have (in units where  $\hbar = k_B = 1$ )

$$\rho_\gamma = \frac{\pi^2}{15} T^4; \quad p_\gamma = \frac{1}{3} \rho_\gamma; \quad s_\gamma = \frac{4\rho_\gamma}{3T}; \quad n_\gamma = \frac{2\zeta(3)}{\pi^2} T^3, \quad (22.39)$$

with  $2\zeta(3)/\pi^2 \simeq 0.2436$ . Note that Eq. (22.10) can be converted into an equation for entropy conservation. Recognizing that  $\dot{p} = s\dot{T}$ , Eq. (22.10) becomes

$$d(sR^3)/dt = 0. \quad (22.40)$$

For radiation, this corresponds to the relationship between expansion and cooling,  $T \propto R^{-1}$  in an adiabatically expanding universe. Note also that both  $s$  and  $n_\gamma$  scale as  $T^3$ .



**Figure 22.3:** The effective numbers of relativistic degrees of freedom as a function of temperature. The sharp drop corresponds to the quark-hadron transition. The solid curve assume a QCD scale of 150 MeV, while the dashed curve assumes 450 MeV.

### 22.3.2. Radiation content of the Early Universe :

At the very high temperatures associated with the early Universe, massive particles are pair produced, and are part of the thermal bath. If for a given particle species  $i$  we have  $T \gg m_i$ , then we can neglect the mass in Eq. (22.34) to Eq. (22.38), and the thermodynamic quantities are easily computed as in Eq. (22.39). In general, we can approximate the energy density (at high temperatures) by including only those particles with  $m_i \ll T$ . In this case, we have

$$\rho = \left( \sum_B g_B + \frac{7}{8} \sum_F g_F \right) \frac{\pi^2}{30} T^4 \equiv \frac{\pi^2}{30} N(T) T^4, \quad (22.41)$$

where  $g_{B(F)}$  is the number of degrees of freedom of each boson (fermion) and the sum runs over all boson and fermion states with  $m \ll T$ . The factor of  $7/8$  is due to the difference between the Fermi and Bose integrals. Eq. (22.41) defines the effective number of degrees of freedom,  $N(T)$ , by taking into account new particle degrees of freedom as the temperature is raised. This quantity is plotted in Fig. 22.3 [38]. For a more recent examination of  $N(T)$  near the QCD transition, see [39].

The value of  $N(T)$  at any given temperature depends on the particle physics model. In the standard  $SU(3) \times SU(2) \times U(1)$  model, we can specify  $N(T)$  up to temperatures of  $O(100)$  GeV. The change in  $N$  (ignoring mass effects) can be seen in the table below.

Temperature	New Particles	$4N(T)$
$T < m_e$	$\gamma$ 's + $\nu$ 's	29
$m_e < T < m_\mu$	$e^\pm$	43
$m_\mu < T < m_\pi$	$\mu^\pm$	57
$m_\pi < T < T_c^\dagger$	$\pi$ 's	69
$T_c < T < m_{\text{strange}}$	$\pi$ 's + $u, \bar{u}, d, \bar{d}$ + gluons	205
$m_s < T < m_{\text{charm}}$	$s, \bar{s}$	247
$m_c < T < m_\tau$	$c, \bar{c}$	289
$m_\tau < T < m_{\text{bottom}}$	$\tau^\pm$	303
$m_b < T < m_{W,Z}$	$b, \bar{b}$	345
$m_{W,Z} < T < m_{\text{Higgs}}$	$W^\pm, Z$	381
$m_H < T < m_{\text{top}}$	$H^0$	385
$m_t < T$	$t, \bar{t}$	427

$^\dagger T_c$  corresponds to the confinement-deconfinement transition between quarks and hadrons.

At higher temperatures,  $N(T)$  will be model-dependent. For example, in the minimal  $SU(5)$  model, one needs to add 24 states to  $N(T)$  for the  $X$  and  $Y$  gauge bosons, another 24 from the adjoint Higgs, and another 6 (in addition to the 4 already counted in  $W^\pm, Z$ , and  $H$ ) from the  $\bar{5}$  of Higgs. Hence for  $T > m_X$  in minimal  $SU(5)$ ,  $N(T) = 160.75$ . In a supersymmetric model this would at least double, with some changes possibly necessary in the table if the lightest supersymmetric particle has a mass below  $m_t$ .

In the radiation-dominated epoch, Eq. (22.10) can be integrated (neglecting the  $T$ -dependence of  $N$ ) giving us a relationship between the age of the Universe and its temperature

$$t = \left( \frac{90}{32\pi^3 G_N N(T)} \right)^{1/2} T^{-2}. \quad (22.42)$$

Put into a more convenient form

$$t T_{\text{MeV}}^2 = 2.4 [N(T)]^{-1/2}, \quad (22.43)$$

where  $t$  is measured in seconds and  $T_{\text{MeV}}$  in units of MeV.

**22.3.3. Neutrinos and equilibrium :** Due to the expansion of the Universe, certain rates may be too slow to either establish or maintain equilibrium. Quantitatively, for each particle  $i$ , as a minimal condition for equilibrium, we will require that some rate  $\Gamma_i$  involving that type be larger than the expansion rate of the Universe or

$$\Gamma_i > H. \quad (22.44)$$

Recalling that the age of the Universe is determined by  $H^{-1}$ , this condition is equivalent to requiring that on average, at least one interaction has occurred over the lifetime of the Universe.

A good example for a process which goes in and out of equilibrium is the weak interactions of neutrinos. On dimensional grounds, one can estimate the thermally averaged scattering cross section:

$$\langle \sigma v \rangle \sim O(10^{-2}) T^2 / m_W^4 \quad (22.45)$$

for  $T \lesssim m_W$ . Recalling that the number density of leptons is  $n \propto T^3$ , we can compare the weak interaction rate,  $\Gamma_{\text{wk}} \sim n \langle \sigma v \rangle$ , with the expansion rate,

$$H = \left( \frac{8\pi G_N \rho}{3} \right)^{1/2} = \left( \frac{8\pi^3}{90} N(T) \right)^{1/2} T^2 / M_P \quad (22.46)$$

$$\sim 1.66 N(T)^{1/2} T^2 / M_P,$$

where the Planck mass  $M_P = G_N^{-1/2} = 1.22 \times 10^{19}$  GeV.

Neutrinos will be in equilibrium when  $\Gamma_{\text{wk}} > H$  or

$$T > (500 m_W^4 / M_P)^{1/3} \sim 1 \text{ MeV}. \quad (22.47)$$

However, this condition assumes  $T \ll m_W$ ; for higher temperatures, we should write  $\langle \sigma v \rangle \sim O(10^{-2}) / T^2$ , so that  $\Gamma \sim 10^{-2} T$ . Thus, in the very early stages of expansion, at temperatures  $T \gtrsim 10^{-2} M_P / \sqrt{N}$ , equilibrium will not have been established.

Having attained a quasi-equilibrium stage, the Universe then cools further to the point where the interaction and expansion timescales match once again. The temperature at which these rates are equal is commonly referred to as the neutrino decoupling or freeze-out temperature and is defined by  $\Gamma_{\text{wk}}(T_d) = H(T_d)$ . For  $T < T_d$ , neutrinos drop out of equilibrium. The Universe becomes transparent to neutrinos and their momenta simply redshift with the cosmic expansion. The effective neutrino temperature will simply fall with  $T \sim 1/R$ .

Soon after decoupling,  $e^\pm$  pairs in the thermal background begin to annihilate (when  $T \lesssim m_e$ ). Because the neutrinos are decoupled, the energy released due to annihilation heats up the photon background relative to the neutrinos. The change in the photon temperature can be easily computed from entropy conservation. The neutrino entropy must be conserved separately from the entropy of interacting particles. A straightforward computation yields

$$T_\nu = (4/11)^{1/3} T_\gamma \simeq 1.9 \text{ K}. \quad (22.48)$$

The total entropy density is therefore given by the contribution from photons and 3 flavors of neutrinos

$$s = \frac{4\pi^2}{30} \left( 2 + \frac{21}{4} (T_\nu / T_\gamma)^3 \right) T_\gamma^3 = \frac{4\pi^2}{30} \left( 2 + \frac{21}{11} \right) T_\gamma^3 = 7.04 n_\gamma. \quad (22.49)$$



Similarly, the total relativistic energy density is given by

$$\rho_r = \frac{\pi^2}{30} \left[ 2 + \frac{21}{4} (T_\nu/T_\gamma)^4 \right] T_\gamma^4 \simeq 1.68 \rho_\gamma. \quad (22.50)$$

In practice, a small correction is needed to this, since neutrinos are not totally decoupled at  $e^\pm$  annihilation: the effective number of massless neutrino species is 3.046, rather than 3 [40].

This expression ignores neutrino rest masses, but current oscillation data require at least one neutrino eigenstate to have a mass exceeding 0.05 eV. In this minimal case,  $\Omega_\nu h^2 = 5 \times 10^{-4}$ , so the neutrino contribution to the matter budget would be negligibly small (which is our normal assumption). However, a nearly degenerate pattern of mass eigenstates could allow larger densities, since oscillation experiments only measure differences in  $m^2$  values. Note that a 0.05-eV neutrino has  $kT_\nu = m_\nu$  at  $z \simeq 297$ , so the above expression for the total present relativistic density is really only an extrapolation. However, neutrinos are almost certainly relativistic at all epochs where the radiation content of the Universe is dynamically significant.

#### 22.3.4. Field Theory and Phase transitions :

It is very likely that the Universe has undergone one or more phase transitions during the course of its evolution [41–44]. Our current vacuum state is described by  $SU(3)_c \times U(1)_{em}$ , which in the Standard Model is a remnant of an unbroken  $SU(3)_c \times SU(2)_L \times U(1)_Y$  gauge symmetry. Symmetry breaking occurs when a non-singlet gauge field (the Higgs field in the Standard Model) picks up a non-vanishing vacuum expectation value, determined by a scalar potential. For example, a simple (non-gauged) potential describing symmetry breaking is  $V(\phi) = \frac{1}{4}\lambda\phi^4 - \frac{1}{2}\mu^2\phi^2 + V(0)$ . The resulting expectation value is simply  $\langle\phi\rangle = \mu/\sqrt{\lambda}$ .

In the early Universe, finite temperature radiative corrections typically add terms to the potential of the form  $\phi^2 T^2$ . Thus, at very high temperatures, the symmetry is restored and  $\langle\phi\rangle = 0$ . As the Universe cools, depending on the details of the potential, symmetry breaking will occur via a first order phase transition in which the field tunnels through a potential barrier, or via a second order transition in which the field evolves smoothly from one state to another (as would be the case for the above example potential).

The evolution of scalar fields can have a profound impact on the early Universe. The equation of motion for a scalar field  $\phi$  can be derived from the energy-momentum tensor

$$T_{\mu\nu} = \partial_\mu\phi\partial_\nu\phi - \frac{1}{2}g_{\mu\nu}\partial_\rho\phi\partial^\rho\phi - g_{\mu\nu}V(\phi). \quad (22.51)$$

By associating  $\rho = T_{00}$  and  $p = R^{-2}(t)T_{ii}$  we have

$$\begin{aligned} \rho &= \frac{1}{2}\dot{\phi}^2 + \frac{1}{2}R^{-2}(t)(\nabla\phi)^2 + V(\phi) \\ p &= \frac{1}{2}\dot{\phi}^2 - \frac{1}{6}R^{-2}(t)(\nabla\phi)^2 - V(\phi), \end{aligned} \quad (22.52)$$

and from Eq. (22.10) we can write the equation of motion (by considering a homogeneous region, we can ignore the gradient terms)

$$\ddot{\phi} + 3H\dot{\phi} = -\partial V/\partial\phi. \quad (22.53)$$

#### 22.3.5. Inflation :

In Sec. 22.2.4, we discussed some of the problems associated with the standard Big-Bang model. However, during a phase transition, our assumptions of an adiabatically expanding universe are generally not valid. If, for example, a phase transition occurred in the early Universe such that the field evolved slowly from the symmetric state to the global minimum, the Universe may have been dominated by the vacuum energy density associated with the potential near  $\phi \simeq 0$ . During this period of slow evolution, the energy density due to radiation will fall below the vacuum energy density,  $\rho \ll V(0)$ . When this happens, the expansion rate will be dominated by the constant  $V(0)$ , and we obtain the exponentially expanding solution given in Eq. (22.20). When the field evolves towards the global minimum it will

begin to oscillate about the minimum, energy will be released during its decay, and a hot thermal universe will be restored. If released fast enough, it will produce radiation at a temperature  $NT_R^4 \lesssim V(0)$ . In this reheating process, entropy has been created and the final value of  $RT$  is greater than the initial value of  $RT$ . Thus, we see that, during a phase transition, the relation  $RT \sim \text{constant}$  need not hold true. This is the basis of the inflationary Universe scenario [45–47].

If, during the phase transition, the value of  $RT$  changed by a factor of  $O(10^{29})$ , the cosmological problems discussed above would be solved. The observed isotropy would be generated by the immense expansion; one small causal region could get blown up, and thus our entire visible Universe would have been in thermal contact some time in the past. In addition, the density parameter  $\Omega$  would have been driven to 1 (with exponential precision). Density perturbations will be stretched by the expansion,  $\lambda \sim R(t)$ . Thus it will appear that  $\lambda \gg H^{-1}$  or that the perturbations have left the horizon, where in fact the size of the causally connected region is now no longer simply  $H^{-1}$ . However, not only does inflation offer an explanation for large scale perturbations, it also offers a source for the perturbations themselves through quantum fluctuations.

Problems with early models of inflation which were based on a either first order [48] or second order [49,50] phase transition of a Grand Unified Theory led models invoking a completely new scalar field: the inflaton,  $\phi$ . The potential of this field,  $V(\phi)$ , needs to have a very low gradient and curvature in order to match observed metric fluctuations. For a more thorough discussion of the problems of early models and a host of current models being studying see the upcoming review on inflation of this *Review*. In most current inflation models, reheated bubbles typically do not percolate, so inflation is ‘eternal’ and continues with exponential expansion in the region outside bubbles. These causally disconnected bubble universes constitute a ‘multiverse’, where low-energy physics can vary between different bubbles. This has led to a controversial ‘anthropic’ approach to cosmology [51–53], where observer selection within the multiverse can be introduced as a means of understanding e.g. why the observed level of vacuum energy is so low (because larger values suppress growth of structure).

#### 22.3.6. Baryogenesis :

The Universe appears to be populated exclusively with matter rather than antimatter. Indeed antimatter is only detected in accelerators or in cosmic rays. However, the presence of antimatter in the latter is understood to be the result of collisions of primary particles in the interstellar medium. There is in fact strong evidence against primary forms of antimatter in the Universe. Furthermore, the density of baryons compared to the density of photons is extremely small,  $\eta \sim 10^{-10}$ .

The production of a net baryon asymmetry requires baryon number violating interactions,  $C$  and  $CP$  violation and a departure from thermal equilibrium [54]. The first two of these ingredients are expected to be contained in grand unified theories as well as in the non-perturbative sector of the Standard Model, the third can be realized in an expanding universe where as we have seen interactions come in and out of equilibrium.

There are several interesting and viable mechanisms for the production of the baryon asymmetry. While, we can not review any of them here in any detail, we mention some of the important scenarios. In all cases, all three ingredients listed above are incorporated. One of the first mechanisms was based on the out of equilibrium decay of a massive particle such as a superheavy GUT gauge of Higgs boson [55,56]. A novel mechanism involving the decay of flat directions in supersymmetric models is known as the Affleck-Dine scenario [57]. There is also the possibility of generating the baryon asymmetry at the electro-weak scale using the non-perturbative interactions of sphalerons [58]. Because these interactions conserve the sum of baryon and lepton number,  $B + L$ , it is possible to first generate a lepton asymmetry (e.g., by the out-of-equilibrium decay of a superheavy right-handed neutrino), which is converted to a baryon asymmetry at the electro-weak scale [59]. This mechanism is known as lepto-baryogenesis.



### 22.3.7. Nucleosynthesis :

An essential element of the standard cosmological model is Big-Bang nucleosynthesis (BBN), the theory which predicts the abundances of the light element isotopes D,  $^3\text{He}$ ,  $^4\text{He}$ , and  $^7\text{Li}$ . Nucleosynthesis takes place at a temperature scale of order 1 MeV. The nuclear processes lead primarily to  $^4\text{He}$ , with a primordial mass fraction of about 25%. Lesser amounts of the other light elements are produced: about  $10^{-5}$  of D and  $^3\text{He}$  and about  $10^{-10}$  of  $^7\text{Li}$  by number relative to H. The abundances of the light elements depend almost solely on one key parameter, the baryon-to-photon ratio,  $\eta$ . The nucleosynthesis predictions can be compared with observational determinations of the abundances of the light elements. Consistency between theory and observations driven primarily by recent D/H measurements [60] leads to a range of

$$5.8 \times 10^{-10} < \eta < 6.6 \times 10^{-10} . \quad (22.54)$$

$\eta$  is related to the fraction of  $\Omega$  contained in baryons,  $\Omega_b$

$$\Omega_b = 3.66 \times 10^7 \eta h^{-2} , \quad (22.55)$$

or  $10^{10} \eta = 274 \Omega_b h^2$ . The Planck result [31] for  $\Omega_b h^2$  of  $0.0223 \pm 0.0002$  translates into a value of  $\eta = 6.09 \pm 0.06$ . This result can be used to ‘predict’ the light element abundance which can in turn be compared with observation [61]. The resulting D/H abundance is in excellent agreement with that found in quasar absorption systems. It is in reasonable agreement with the helium abundance observed in extra-galactic HII regions (once systematic uncertainties are accounted for), but is in poor agreement with the Li abundance observed in the atmospheres of halo dwarf stars [62]. (See the review on BBN—Sec. 24 of this *Review* for a detailed discussion of BBN or references [63,64,65].)

### 22.3.8. The transition to a matter-dominated Universe :

In the Standard Model, the temperature (or redshift) at which the Universe undergoes a transition from a radiation dominated to a matter dominated Universe is determined by the amount of dark matter. Assuming three nearly massless neutrinos, the energy density in radiation at temperatures  $T \ll 1$  MeV, is given by

$$\rho_r = \frac{\pi^2}{30} \left[ 2 + \frac{21}{4} \left( \frac{4}{11} \right)^{4/3} \right] T^4 . \quad (22.56)$$

In the absence of non-baryonic dark matter, the matter density can be written as

$$\rho_m = m_N \eta n_\gamma , \quad (22.57)$$

where  $m_N$  is the nucleon mass. Recalling that  $n_\gamma \propto T^3$  [cf. Eq. (22.39)], we can solve for the temperature or redshift at the matter-radiation equality when  $\rho_r = \rho_m$ ,

$$T_{\text{eq}} = 0.22 m_N \eta \quad \text{or} \quad (1 + z_{\text{eq}}) = 0.22 \eta \frac{m_N}{T_0} , \quad (22.58)$$

where  $T_0$  is the present temperature of the microwave background. For  $\eta = 6.1 \times 10^{-10}$ , this corresponds to a temperature  $T_{\text{eq}} \simeq 0.13$  eV or  $(1 + z_{\text{eq}}) \simeq 550$ . A transition this late is very problematic for structure formation (see Sec. 22.4.5).

The redshift of matter domination can be pushed back significantly if non-baryonic dark matter is present. If instead of Eq. (22.57), we write

$$\rho_m = \Omega_m \rho_c \left( \frac{T}{T_0} \right)^3 , \quad (22.59)$$

we find that

$$T_{\text{eq}} = 0.9 \frac{\Omega_m \rho_c}{T_0^3} \quad \text{or} \quad (1 + z_{\text{eq}}) = 2.4 \times 10^4 \Omega_m h^2 . \quad (22.60)$$

## 22.4. The Universe at late times

### 22.4.1. The CMB :

One form of the infamous Olbers’ paradox says that, in Euclidean space, surface brightness is independent of distance. Every line of sight will terminate on matter that is hot enough to be ionized and so scatter photons:  $T \gtrsim 10^3$  K; the sky should therefore shine as brightly as the surface of the Sun. The reason the night sky is dark is entirely due to the expansion, which cools the radiation temperature to 2.73 K. This gives a Planck function peaking at around 1 mm to produce the microwave background (CMB).

The CMB spectrum is a very accurate match to a Planck function [66]. (See the review on CBR—Sec. 28 of this *Review*.) The COBE estimate of the temperature is [67]

$$T = 2.7255 \pm 0.0006 \text{ K} . \quad (22.61)$$

The lack of any distortion of the Planck spectrum is a strong physical constraint. It is very difficult to account for in any expanding universe other than one that passes through a hot stage. Alternative schemes for generating the radiation, such as thermalization of starlight by dust grains, inevitably generate a superposition of temperatures. What is required in addition to thermal equilibrium is that  $T \propto 1/R$ , so that radiation from different parts of space appears identical.

Although it is common to speak of the CMB as originating at ‘recombination’, a more accurate terminology is the era of ‘last scattering’. In practice, this takes place at  $z \simeq 1100$ , almost independently of the main cosmological parameters, at which time the fractional ionization is very small. This occurred when the age of the Universe was about 400,000 years. (See the review on CBR—Sec. 28 of this *Review* for a full discussion of the CMB.)

### 22.4.2. Matter in the Universe :

One of the main tasks of cosmology is to measure the density of the Universe, and how this is divided between dark matter and baryons. The baryons consist partly of stars, with  $0.002 \lesssim \Omega_* \lesssim 0.003$  [68] but mainly inhabit the intergalactic medium (IGM). One powerful way in which this can be studied is via the absorption of light from distant luminous objects such as quasars. Even very small amounts of neutral hydrogen can absorb rest-frame UV photons (the Gunn-Peterson effect), and should suppress the continuum by a factor  $\exp(-\tau)$ , where

$$\tau \simeq 10^{4.62} h^{-1} \left[ \frac{n_{\text{HI}}(z)/\text{m}^{-3}}{(1+z)\sqrt{1+\Omega_m z}} \right] , \quad (22.62)$$

and this expression applies while the Universe is matter dominated ( $z \gtrsim 1$  in the  $\Omega_m = 0.3$   $\Omega_v = 0.7$  model). It is possible that this general absorption has now been seen at  $z = 6.2 - 6.4$  [69]. At lower redshifts, the dominant effect on the spectrum is a ‘forest’ of narrow absorption lines, which produce a mean  $\tau = 1$  in the Ly $\alpha$  forest at about  $z = 3$ , and so we have  $\Omega_{\text{HI}} \simeq 10^{-6.7} h^{-1}$ . This is such a small number that clearly the IGM is very highly ionized at these redshifts.

The Ly $\alpha$  forest is of great importance in pinning down the abundance of deuterium. Because electrons in deuterium differ in reduced mass by about 1 part in 4000 compared to hydrogen, each absorption system in the Ly $\alpha$  forest is accompanied by an offset deuterium line. By careful selection of systems with an optimal HI column density, a measurement of the D/H ratio can be made. This has now been done with high accuracy in 5 quasars, with consistent results [60]. Combining these determinations with the theory of primordial nucleosynthesis yields a baryon density of  $\Omega_b h^2 = 0.021 - 0.023$  (95% confidence) in excellent agreement with the Planck result. (See also the review on BBN—Sec. 24 of this *Review*.)

Ionized IGM can also be detected in emission when it is densely clumped, via bremsstrahlung radiation. This generates the spectacular X-ray emission from rich clusters of galaxies. Studies of this phenomenon allow us to achieve an accounting of the total baryonic material in clusters. Within the central  $\simeq 1$  Mpc, the masses in stars, X-ray emitting gas and total dark matter can be determined with reasonable accuracy (perhaps 20% rms), and this allows a minimum baryon fraction to be determined [70,71]:

$$\frac{M_{\text{baryons}}}{M_{\text{total}}} \gtrsim 0.009 + (0.066 \pm 0.003) h^{-3/2} . \quad (22.63)$$

Because clusters are the largest collapsed structures, it is reasonable to take this as applying to the Universe as a whole. This equation implies a minimum baryon fraction of perhaps 12% (for reasonable  $h$ ), which is too high for  $\Omega_m = 1$  if we take  $\Omega_b h^2 \simeq 0.02$  from nucleosynthesis. This is therefore one of the more robust arguments in favor of  $\Omega_m \simeq 0.3$ . (See the review on Cosmological Parameters—Sec. 25 of this *Review*.) This argument is also consistent with the inference on  $\Omega_m$  that can be made from Fig. 22.2.

This method is much more robust than the older classical technique for weighing the Universe: ‘ $L \times M/L$ ’. The overall light density of the Universe is reasonably well determined from redshift surveys of galaxies, so that a good determination of mass  $M$  and luminosity  $L$  for a single object suffices to determine  $\Omega_m$  if the mass-to-light ratio is universal.

#### 22.4.3. Gravitational lensing :

A robust method for determining masses in cosmology is to use gravitational light deflection. Most systems can be treated as a geometrically thin gravitational lens, where the light bending is assumed to take place only at a single distance. Simple geometry then determines a mapping between the coordinates in the intrinsic source plane and the observed image plane:

$$\alpha(D_L \theta_1) = \frac{D_S}{D_{LS}}(\theta_1 - \theta_S), \quad (22.64)$$

where the angles  $\theta_1, \theta_S$  and  $\alpha$  are in general two-dimensional vectors on the sky. The distances  $D_{LS}$  etc. are given by an extension of the usual distance-redshift formula:

$$D_{LS} = \frac{R_0 S_k(\chi_S - \chi_L)}{1 + z_S}. \quad (22.65)$$

This is the angular-diameter distance for objects on the source plane as perceived by an observer on the lens.

Solutions of this equation divide into weak lensing, where the mapping between source plane and image plane is one-to-one, and strong lensing, in which multiple imaging is possible. For circularly-symmetric lenses, an on-axis source is multiply imaged into a ‘caustic’ ring, whose radius is the Einstein radius:

$$\begin{aligned} \theta_E &= \left( 4GM \frac{D_{LS}}{D_L D_S} \right)^{1/2} \\ &= \left( \frac{M}{10^{11.09} M_\odot} \right)^{1/2} \left( \frac{D_L D_S / D_{LS}}{\text{Gpc}} \right)^{-1/2} \text{ arcsec}. \end{aligned} \quad (22.66)$$

The observation of ‘arcs’ (segments of near-perfect Einstein rings) in rich clusters of galaxies has thus given very accurate masses for the central parts of clusters—generally in good agreement with other indicators, such as analysis of X-ray emission from the cluster IGM [72].

Gravitational lensing has also developed into a particularly promising probe of cosmological structure on 10 to 100 Mpc scales. Weak image distortions manifest themselves as an additional ellipticity of galaxy images (‘shear’), which can be observed by averaging many images together (the corresponding flux amplification is less readily detected). The result is a ‘cosmic shear’ field of order 1% ellipticity, coherent over scales of around 30 arcmin, which is directly related to the cosmic mass field, without any astrophysical uncertainties. For this reason, weak lensing is seen as potentially the cleanest probe of matter fluctuations, next to the CMB. Already, impressive results have been obtained in measuring cosmological parameters, based on survey data from only  $\sim 150 \text{ deg}^2$  [73]. A particular strength of lensing is its ability to measure the amplitude of mass fluctuations; this can be deduced from the amplitude of CMB fluctuations, but only with low precision on account of the poorly-known optical depth due to Compton scattering after reionization. However, the effect of weak lensing on the CMB map itself can be detected via the induced non-Gaussian signal, and this gives the CMB greater internal power [74].

#### 22.4.4. Density Fluctuations :

The overall properties of the Universe are very close to being homogeneous; and yet telescopes reveal a wealth of detail on scales varying from single galaxies to large-scale structures of size exceeding 100 Mpc. The existence of these structures must be telling us something important about the initial conditions of the Big Bang, and about the physical processes that have operated subsequently. This motivates the study of the density perturbation field, defined as

$$\delta(\mathbf{x}) \equiv \frac{\rho(\mathbf{x}) - \langle \rho \rangle}{\langle \rho \rangle}. \quad (22.67)$$

A critical feature of the  $\delta$  field is that it inhabits a universe that is isotropic and homogeneous in its large-scale properties. This suggests that the statistical properties of  $\delta$  should also be statistically homogeneous—*i.e.*, it is a stationary random process.

It is often convenient to describe  $\delta$  as a Fourier superposition:

$$\delta(\mathbf{x}) = \sum \delta_{\mathbf{k}} e^{-i\mathbf{k} \cdot \mathbf{x}}. \quad (22.68)$$

We avoid difficulties with an infinite universe by applying periodic boundary conditions in a cube of some large volume  $V$ . The cross-terms vanish when we compute the variance in the field, which is just a sum over modes of the power spectrum

$$\langle \delta^2 \rangle = \sum |\delta_{\mathbf{k}}|^2 \equiv \sum P(k). \quad (22.69)$$

Note that the statistical nature of the fluctuations must be isotropic, so we write  $P(k)$  rather than  $P(\mathbf{k})$ . The  $\langle \dots \rangle$  average here is a volume average. Cosmological density fields are an example of an ergodic process, in which the average over a large volume tends to the same answer as the average over a statistical ensemble.

The statistical properties of discrete objects sampled from the density field are often described in terms of  $N$ -point correlation functions, which represent the excess probability over random for finding one particle in each of  $N$  boxes in a given configuration. For the 2-point case, the correlation function is readily shown to be identical to the autocorrelation function of the  $\delta$  field:  $\xi(r) = \langle \delta(x)\delta(x+r) \rangle$ .

The power spectrum and correlation function are Fourier conjugates, and thus are equivalent descriptions of the density field (similarly,  $k$ -space equivalents exist for the higher-order correlations). It is convenient to take the limit  $V \rightarrow \infty$  and use  $k$ -space integrals, defining a dimensionless power spectrum, which measures the contribution to the fractional variance in density per unit logarithmic range of scale, as  $\Delta^2(k) = d\langle \delta^2 \rangle / d \ln k = V k^3 P(k) / 2\pi^2$ :

$$\xi(r) = \int \Delta^2(k) \frac{\sin kr}{kr} d \ln k; \quad \Delta^2(k) = \frac{2}{\pi} k^3 \int_0^\infty \xi(r) \frac{\sin kr}{kr} r^2 dr. \quad (22.70)$$

For many years, an adequate approximation to observational data on galaxies was  $\xi = (r/r_0)^{-\gamma}$ , with  $\gamma \simeq 1.8$  and  $r_0 \simeq 5 h^{-1} \text{ Mpc}$ . Modern surveys are now able to probe into the large-scale linear regime where unaltered traces of the curved post-recombination spectrum can be detected [75–77].

#### 22.4.5. Formation of cosmological structure :

The simplest model for the generation of cosmological structure is gravitational instability acting on some small initial fluctuations (for the origin of which a theory such as inflation is required). If the perturbations are adiabatic (*i.e.*, fractionally perturb number densities of photons and matter equally), the linear growth law for matter perturbations is simple:

$$\delta \propto \begin{cases} a^2(t) & (\text{radiation domination; } \Omega_r = 1) \\ a(t) & (\text{matter domination; } \Omega_m = 1) \end{cases} \quad (22.71)$$

For low-density universes, the growth is slower:

$$d \ln \delta / d \ln a \simeq \Omega_m^{-\gamma}(a), \quad (22.72)$$

where the parameter  $\gamma$  is close to 0.55 independent of the vacuum density [78].

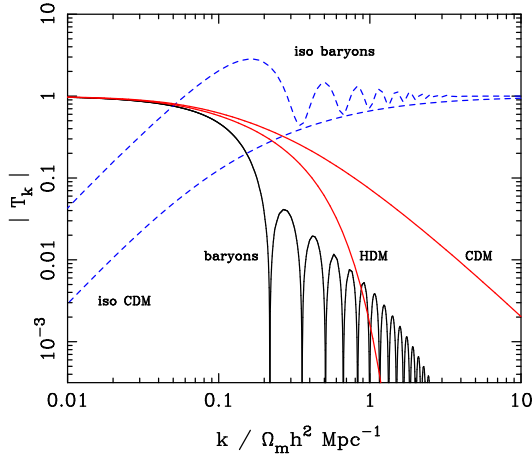
The alternative perturbation mode is isocurvature: only the equation of state changes, and the total density is initially unperturbed. These modes perturb the total entropy density, and thus induce additional large-scale CMB anisotropies [79]. Although the character of perturbations in the simplest inflationary theories are purely adiabatic, correlated adiabatic and isocurvature modes are predicted in many models; the simplest example is the curvaton, which is a scalar field that decays to yield a perturbed radiation density. If the matter content already exists at this time, the overall perturbation field will have a significant isocurvature component. Such a prediction is inconsistent with current CMB data [80], and most analyses of CMB and large scale structure (LSS) data assume the adiabatic case to hold exactly.

Linear evolution preserves the shape of the power spectrum. However, a variety of processes mean that growth actually depends on the matter content:

- (1) Pressure opposes gravity effectively for wavelengths below the horizon length while the Universe is radiation dominated. The *comoving* horizon size at  $z_{\text{eq}}$  is therefore an important scale:

$$D_H(z_{\text{eq}}) = \frac{2(\sqrt{2}-1)}{(\Omega_m z_{\text{eq}})^{1/2} H_0} = \frac{16.0}{\Omega_m h^2} \text{Mpc}. \quad (22.73)$$

- (2) At early times, dark matter particles will undergo free streaming at the speed of light, and so erase all scales up to the horizon—a process that only ceases when the particles go nonrelativistic. For light massive neutrinos, this happens at  $z_{\text{eq}}$ ; all structure up to the horizon-scale power-spectrum break is in fact erased. Hot(cold) dark matter models are thus sometimes dubbed large(small)-scale damping models.
- (3) A further important scale arises where photon diffusion can erase perturbations in the matter–radiation fluid; this process is named Silk damping.



**Figure 22.4:** A plot of transfer functions for various models. For adiabatic models,  $T_k \rightarrow 1$  at small  $k$ , whereas the opposite is true for isocurvature models. For dark-matter models, the characteristic wavenumber scales proportional to  $\Omega_m h^2$ . The scaling for baryonic models does not obey this exactly; the plotted cases correspond to  $\Omega_m = 1$ ,  $h = 0.5$ .

The overall effect is encapsulated in the transfer function, which gives the ratio of the late-time amplitude of a mode to its initial value (see Fig. 22.4). The overall power spectrum is thus the primordial scalar-mode power law, times the square of the transfer function:

$$P(k) \propto k^{n_s} T_k^2. \quad (22.74)$$

The most generic power-law index is  $n_s = 1$ : the ‘Zeldovich’ or ‘scale-invariant’ spectrum. Inflationary models tend to predict a small ‘tilt’:  $|n_s - 1| \lesssim 0.03$  [12,13]. On the assumption that the dark matter is cold, the power spectrum then depends on 5 parameters:  $n_s$ ,  $h$ ,  $\Omega_b$ ,  $\Omega_c$  ( $\equiv \Omega_m - \Omega_b$ ) and an overall amplitude. The latter is

often specified as  $\sigma_8$ , the linear-theory fractional rms in density when a spherical filter of radius  $8 h^{-1} \text{Mpc}$  is applied in linear theory. This scale can be probed directly via weak gravitational lensing, and also via its effect on the abundance of rich galaxy clusters. The favored value from the latter is approximately [81]

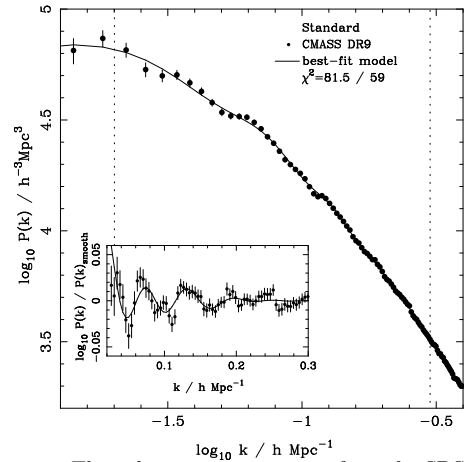
$$\sigma_8 \simeq [0.813 \pm 0.013 (\text{stat}) \pm 0.024 (\text{sys})] (\Omega_m/0.25)^{-0.47}, \quad (22.75)$$

which is consistent with the Planck values of  $(\sigma_8, \Omega_m) = (0.815 \pm 0.009, 0.308 \pm 0.012)$ .

A direct measure of mass inhomogeneity is valuable, since the galaxies inevitably are biased with respect to the mass. This means that the fractional fluctuations in galaxy number,  $\delta n/n$ , may differ from the mass fluctuations,  $\delta \rho/\rho$ . It is commonly assumed that the two fields obey some proportionality on large scales where the fluctuations are small,  $\delta n/n = b \delta \rho/\rho$ , but even this is not guaranteed [82].

The main shape of the transfer function is a break around the horizon scale at  $z_{\text{eq}}$ , which depends just on  $\Omega_m h$  when wavenumbers are measured in observable units ( $h \text{Mpc}^{-1}$ ). For reasonable baryon content, weak oscillations in the transfer function are also expected, and these BAOs (Baryon Acoustic Oscillations) have been clearly detected [83,84]. As well as directly measuring the baryon fraction, the scale of the oscillations directly measures the acoustic horizon at decoupling; this can be used as an additional standard ruler for cosmological tests, and the BAO signature has become one of the most important applications of large galaxy surveys. Overall, current power-spectrum data [75–77] favor  $\Omega_m h \simeq 0.20$  and a baryon fraction of about 0.15 for  $n_s = 1$  (see Fig. 22.5).

In principle, accurate data over a wide range of  $k$  could determine both  $\Omega_m h$  and  $n_s$ , but in practice there is a strong degeneracy between these. In order to constrain  $n_s$  itself, it is necessary to examine data on anisotropies in the CMB.



**Figure 22.5:** The galaxy power spectrum from the SDSS BOSS survey [77]. The solid points with error bars show the power estimate. The solid line shows a standard  $\Lambda\text{CDM}$  model with  $\Omega_b h^2 \simeq 0.02$  and  $\Omega_m h \simeq 0.2$ . The inset amplifies the region where BAO features are visible. The fact that these perturb the power by  $\sim 20\%$  rather than order unity is direct evidence that the matter content of the universe is dominated by collisionless dark matter.

#### 22.4.6. CMB anisotropies :

The CMB has a clear dipole anisotropy, of magnitude  $1.23 \times 10^{-3}$ . This is interpreted as being due to the Earth’s motion, which is equivalent to a peculiar velocity for the Milky Way of

$$v_{\text{MW}} \simeq 600 \text{ km s}^{-1} \text{ towards } (\ell, b) \simeq (270^\circ, 30^\circ). \quad (22.76)$$

All higher-order multipole moments of the CMB are however much smaller (of order  $10^{-5}$ ), and interpreted as signatures of density fluctuations at last scattering ( $\simeq 1100$ ). To analyze these, the sky is expanded in spherical harmonics as explained in the review on

CBR—Sec. 28 of this *Review*. The dimensionless power per  $\ln k$  or ‘bandpower’ for the CMB is defined as

$$\mathcal{T}^2(\ell) = \frac{\ell(\ell+1)}{2\pi} C_\ell. \quad (22.77)$$

This function encodes information from the three distinct mechanisms that cause CMB anisotropies:

- (1) Gravitational (Sachs–Wolfe) perturbations. Photons from high-density regions at last scattering have to climb out of potential wells, and are thus redshifted.
- (2) Intrinsic (adiabatic) perturbations. In high-density regions, the coupling of matter and radiation can compress the radiation also, giving a higher temperature.
- (3) Velocity (Doppler) perturbations. The plasma has a non-zero velocity at recombination, which leads to Doppler shifts in frequency and hence shifts in brightness temperature.

Because the potential fluctuations obey Poisson’s equation,  $\nabla^2\Phi = 4\pi G\rho\delta$ , and the velocity field satisfies the continuity equation  $\nabla \cdot \mathbf{u} = -\delta$ , the resulting different powers of  $k$  ensure that the Sachs–Wolfe effect dominates on large scales and adiabatic effects on small scales.

The relation between angle and comoving distance on the last-scattering sphere requires the comoving angular-diameter distance to the last-scattering sphere; because of its high redshift, this is effectively identical to the horizon size at the present epoch,  $D_H$ :

$$\begin{aligned} D_H &= \frac{2}{\Omega_m H_0} \quad (\Omega_v = 0) \\ D_H &\simeq \frac{2}{\Omega_m^{0.4} H_0} \quad (\text{flat} : \Omega_m + \Omega_v = 1). \end{aligned} \quad (22.78)$$

These relations show how the CMB is strongly sensitive to curvature: the horizon length at last scattering is  $\propto 1/\sqrt{\Omega_m}$ , so that this subtends an angle that is virtually independent of  $\Omega_m$  for a flat model. Observations of a peak in the CMB power spectrum at relatively large scales ( $\ell \simeq 225$ ) are thus strongly inconsistent with zero- $\Lambda$  models with low density: current CMB + BAO + lensing data require  $\Omega_m + \Omega_v = 1.000 \pm 0.005$  (95%) [31]. (See *e.g.*, Fig. 22.2). This result is unchanged when SN data and the prior on  $H_0$  are included.

In addition to curvature, the CMB encodes information about several other key cosmological parameters. Within the compass of simple adiabatic CDM models, there are 9 of these:

$$\omega_c, \omega_b, \Omega_{\text{tot}}, h, \tau, n_s, n_t, r, Q. \quad (22.79)$$

The symbol  $\omega$  denotes the physical density,  $\Omega h^2$ : the transfer function depends only on the densities of CDM ( $\omega_c$ ) and baryons ( $\omega_b$ ). Transcribing the power spectrum at last scattering into an angular power spectrum brings in the total density parameter ( $\Omega_{\text{tot}} \equiv \Omega_m + \Omega_v = \Omega_c + \Omega_b + \Omega_v$ ) and  $h$ : there is an exact geometrical degeneracy [85] between these that keeps the angular-diameter distance to last scattering invariant, so that models with substantial spatial curvature and large vacuum energy cannot be ruled out without prior knowledge of the Hubble parameter. Alternatively, the CMB alone cannot measure the Hubble parameter.

A further possible degeneracy involves the tensor contribution to the CMB anisotropies. These are important at large scales (up to the horizon scales); for smaller scales, only scalar fluctuations (density perturbations) are important. Each of these components is characterized by a spectral index,  $n$ , and a ratio between the power spectra of tensors and scalars ( $r$ ). See the review on Cosmological Parameters—Sec. 25 of this *Review* for a technical definition of the  $r$  parameter. Finally, the overall amplitude of the spectrum must be specified ( $Q$ ), together with the optical depth to Compton scattering owing to recent reionization ( $\tau$ ). Adding a large tensor contribution reduces the contrast between low  $\ell$  and the peak at  $\ell \simeq 225$  (because the tensor spectrum has no acoustic component). The previous relative height of the peak can be recovered by increasing  $n_s$  to increase the small-scale power in the scalar component; this in turn over-predicts the power at  $\ell \sim 1000$ , but this effect can be

counteracted by raising the baryon density [86]. This approximate 3-way degeneracy is broken as we increase the range of multipoles sampled.

The reason the tensor component is introduced, and why it is so important, is that it is the only non-generic prediction of inflation. Slow-roll models of inflation involve two dimensionless parameters:

$$\epsilon \equiv \frac{M_P^2}{16\pi} \left( \frac{V'}{V} \right)^2, \quad \eta \equiv \frac{M_P^2}{8\pi} \left( \frac{V''}{V} \right), \quad (22.80)$$

where  $V$  is the inflaton potential, and dashes denote derivatives with respect to the inflation field. In terms of these, the tensor-to-scalar ratio is  $r \simeq 16\epsilon$ , and the spectral indices are  $n_s = 1 - 6\epsilon + 2\eta$  and  $n_t = -2\epsilon$ . The natural expectation of inflation is that the quasi-exponential phase ends once the slow-roll parameters become significantly non-zero, so that both  $n_s \neq 1$  and a significant tensor component are expected. These prediction can be avoided in some models, but it is undeniable that observation of such features would be a great triumph for inflation. Cosmology therefore stands at a fascinating point given that the most recent CMB data appear to reject the zero-tensor  $n_s = 1$  model at almost  $6\sigma$ :  $n_s = 0.968 \pm 0.006$  [31]. This rejection is strong enough that it is also able to break the tensor degeneracy, so that no model with  $n_s = 1$  is acceptable, whatever the value of  $r$ .

The current limit on  $r$  is  $< 0.11$  at 95% confidence [87]. In conjunction with the measured value of  $n_s$ , this upper limit sits close to the prediction of a linear potential (i.e.  $|\eta| \ll |\epsilon|$ ). Any further reduction in the limit on  $r$  will force  $\eta$  to be negative – i.e. a convex potential at the point where LSS scales were generated (sometimes called a ‘hilltop’), in contrast to simple early models such as  $V(\phi) = m^2\phi^2$  or  $\lambda\phi^4$ . Examples of models which are currently in excellent agreement with the Planck results are the Starobinsky model of  $\mathcal{R} + \mathcal{R}^2$  gravity [88], or the Higgs-inflation model where the Higgs field is non-minimally coupled [89]. Assuming 55 e-foldings of inflation, these models predict  $n_s = 0.965$  and  $r = 0.0035$ . Assuming that no systematic error in the CMB data can be identified, cosmology has passed a critical hurdle in rejecting scale-invariant fluctuations. The years ahead will be devoted to the task of searching for the tensor fluctuations – for which the main tool will be the polarization of the CMB [14].

#### 22.4.6.1. CMB foregrounds:

As the quality of CMB data improves, there is a growing interest in effects that arise along the line of sight. The CMB temperature is perturbed by dark-matter structures and by Compton scattering from ionized gas. In the former case, we have the Integrated Sachs–Wolfe effect, which is sensitive to the time derivative of the gravitational potential. In the linear regime, this is damped when the universe becomes  $\Lambda$ -dominated, and this is an independent way of detecting  $\Lambda$  [90]. The potential also causes gravitational lensing of the CMB: structures at  $z \sim 1 - 2$  displace features on the CMB sky by about 2 arcmin over coherent degree-scale patches. Detection of these distortions allows a map to be made of overdensity projected from  $z = 0$  to 1100 [74]. This is a very powerful calibration for direct studies of gravitational lensing using galaxies. Finally, Comptonization affects the CMB in two ways: the thermal Sunyaev–Zeldovich effect measures the blurring of photon energies by hot gas; the kinetic Sunyaev–Zeldovich effect is sensitive to the bulk velocity of the gas. Both these effects start to dominate over the intrinsic CMB fluctuations at multipoles  $\ell \gtrsim 2000$  [91].

#### 22.4.7. Probing dark energy and the nature of gravity :

The most radical element of our current cosmological model is the dark energy that accelerates the expansion. The energy density of this component is approximately  $(2.2 \text{ meV})^4$  (for  $w = -1$ ,  $\Omega_v = 0.68$ ,  $h = 0.67$ ), or roughly  $10^{-123} M_P^4$ , and such an un-naturally small number is hard to understand. Various quantum effects (most simply zero-point energy) should make contributions to the vacuum energy density: these may be truncated by new physics at high energy, but this presumably occurs at  $> 1 \text{ TeV}$  scales, not meV; thus the apparent energy scale of the vacuum is at least  $10^{15}$  times smaller than its

natural value. This situation is well analysed in [51], which lists extreme escape routes – especially the multiverse viewpoint, according to which low values of  $\Lambda$  are rare, but high values suppress the formation of structure and observers. It is certainly impressive that Weinberg used such reasoning to predict the value of  $\Lambda$  before any data strongly indicated a non-zero value.

But it may be that the phenomenon of dark energy is entirely illusory. The necessity for this constituent arises from using the Friedmann equation to describe the evolution of the cosmic expansion; if this equation is incorrect, it would require the replacement of Einstein's relativistic theory of gravity with some new alternative. A frontier of current cosmological research is to distinguish these possibilities [92,93]. We also note that it has been suggested that dark energy might be an illusion even within general relativity, owing to an incorrect treatment of averaging in an inhomogeneous Universe [94,95]. Many would argue that a standard Newtonian treatment of such issues should be adequate inside the cosmological horizon, but debate on this issue continues.

Dark Energy can differ from a classical cosmological constant in being a dynamical phenomenon [96,97], *e.g.*, a rolling scalar field (sometimes dubbed 'quintessence'). Empirically, this means that it is endowed with two thermodynamic properties that astronomers can try to measure: the bulk equation of state and the sound speed. If the sound speed is close to the speed of light, the effect of this property is confined to very large scales, and mainly manifests itself in the large-angle multipoles of the CMB anisotropies [98]. The equation of state parameter governs the rate of change of the vacuum density:  $d \ln \rho_v / d \ln a = -3(1+w)$ , so it can be accessed via the evolving expansion rate,  $H(a)$ . This can be measured most cleanly by using the inbuilt natural ruler of large-scale structure: the Baryon Acoustic Oscillation horizon scale [99]:

$$D_{\text{BAO}} \simeq 147 (\Omega_m h^2 / 0.13)^{-0.25} (\Omega_b h^2 / 0.023)^{-0.08} \text{ Mpc}. \quad (22.81)$$

$H(a)$  is measured by radial clustering, since  $dr/dz = c/H$ ; clustering in the plane of the sky measures the integral of this. The expansion rate is also measured by the growth of density fluctuations, where the pressure-free growth equation for the density perturbation is  $\ddot{\delta} + 2H(a)\dot{\delta} = 4\pi G \rho_0 \delta$ . Thus, both the scale and amplitude of density fluctuations are sensitive to  $w(a)$  – but only weakly. These observables change by only typically 0.2% for a 1% change in  $w$ . Current constraints [31] place a constant  $w$  to within 5–10% of  $-1$ , depending on the data combination chosen. A substantial improvement in this precision will require us to limit systematics in data to a few parts in 1000.

Testing whether theories of gravity require revision can also be done using data on cosmological inhomogeneities. Two separate issues arise, concerning the metric perturbation potentials  $\Psi$  and  $\Phi$ , which affect respectively the time and space parts of the metric. In Einstein gravity, these potentials are both equal to the Newtonian gravitational potential, which satisfies Poisson's equation:  $\nabla^2 \Phi / a^2 = 4\pi G \bar{\rho} \delta$ . Empirically, modifications of gravity require us to explore a change with scale and with time of the 'slip' ( $\Psi/\Phi$ ) and the effective  $G$  on the rhs of the Poisson equation. The former aspect can only be probed via gravitational lensing, whereas the latter can be addressed on 10–100 Mpc scales via the growth of clustering. Various schemes for parameterising modified gravity exist, but a practical approach is to assume that the growth rate can be tied to the density parameter:  $d \ln \delta / d \ln a = \Omega_m(a)^\gamma$  [78]. The parameter  $\gamma$  is close to 0.55 for standard relativistic gravity, but can differ by around 0.1 from this value in many non-standard models. Clearly this parameterization is incomplete, since it explicitly rejects the possibility of early dark energy ( $\Omega_m(a) \rightarrow 1$  as  $a \rightarrow 0$ ), but it is a convenient way of capturing the power of various experiments. Current data are consistent with standard  $\Lambda$ CDM [100], and exclude variations in slip or effective  $G$  of larger than a few times 10%.

Current planning envisages a set of satellite probes that, a decade hence, will pursue these fundamental tests via gravitational lensing measurements over thousands of square degrees,  $> 10^8$  redshifts, and photometry of  $> 1000$  supernovae (WFIRST in the USA, Euclid in Europe) [22,23]. These experiments will measure both  $w$  and

the perturbation growth rate to an accuracy of around 1%. The outcome will be either a validation of the standard relativistic vacuum-dominated big bang cosmology at a level of precision far beyond anything attempted to date, or the opening of entirely new directions in cosmological models. For a more complete discussion of dark energy and future probes see the review on Dark Energy—Sec. 27

## References:

1. V.M. Slipher, *Pop. Astr.* **23**, 21 (1915).
2. K. Lundmark, *MNRAS* **84**, 747 (1924).
3. E. Hubble and M.L. Humason, *Astrophys. J.* **74**, 43 (1931).
4. G. Gamow, *Phys. Rev.* **70**, 572 (1946).
5. R.A. Alpher *et al.*, *Phys. Rev.* **73**, 803 (1948).
6. R.A. Alpher and R.C. Herman, *Phys. Rev.* **74**, 1737 (1948).
7. R.A. Alpher and R.C. Herman, *Phys. Rev.* **75**, 1089 (1949).
8. A.A. Penzias and R.W. Wilson, *Astrophys. J.* **142**, 419 (1965).
9. P.J.E. Peebles, *Principles of Physical Cosmology*, Princeton University Press (1993).
10. G. Börner, *The Early Universe: Facts and Fiction*, Springer-Verlag (1988).
11. E.W. Kolb and M.S. Turner, *The Early Universe*, Addison-Wesley (1990).
12. J.A. Peacock, *Cosmological Physics*, Cambridge Univ. Press (1999).
13. A.R. Liddle and D. Lyth, *Cosmological Inflation and Large-Scale Structure*, Cambridge University Press (2000).
14. S. Dodelson, *Modern Cosmology*, Academic Press (2003).
15. V. Mukhanov, *Physical Foundations of Cosmology*, Cambridge University Press (2005).
16. S. Weinberg, *Cosmology*, Oxford Press (2008).
17. E.B. Gliner, *Sov. Phys. JETP* **22**, 378 (1966).
18. Y.B. Zeldovich, (1967), *Sov. Phys. Usp.* **11**, 381 (1968).
19. P.M. Garnavich *et al.*, *Astrophys. J.* **507**, 74 (1998).
20. S. Perlmutter *et al.*, *Phys. Rev. Lett.* **83**, 670 (1999).
21. I. Maor *et al.*, *Phys. Rev. D* **65**, 123003 (2002).
22. A. Albrecht *et al.*, *astro-ph/0609591*.
23. J. Peacock *et al.*, *astro-ph/0610906*.
24. E. Auberg *et al.*, *arXiv:1411.1074*.
25. A.G. Riess *et al.*, *Astrophys. J.* **116**, 1009 (1998).
26. S. Perlmutter *et al.*, *Astrophys. J.* **517**, 565 (1999).
27. A.G. Riess, *Pub. Astron. Soc. Pac.* **112**, 1284 (2000).
28. J.L. Tonry *et al.*, *Astrophys. J.* **594**, 1 (2003).
29. R. Amanullah *et al.*, *Astrophys. J.* **716**, 712 (2010).
30. M. Betoul *et al.*, *Astron. & Astrophys.* **568**, A22 (2014).
31. Planck Collab. 2015 Results XIII, *arXiv:1502.01589*.
32. L. Humphreys *et al.*, *Astrophys. J.* **775**, 13 (2013).
33. P. Astier *et al.*, *Astron. & Astrophys.* **447**, 31 (2006).
34. J.A. Johnson and M. Bolte, *Astrophys. J.* **554**, 888 (2001).
35. R. Cayrel *et al.*, *Nature* **409**, 691 (2001).
36. R. Jimenez and P. Padoan, *Astrophys. J.* **498**, 704 (1998).
37. E. Carretta *et al.*, *Astrophys. J.* **533**, 215 (2000).
38. M. Srednicki *et al.*, *Nucl. Phys.* **B310**, 693 (1988).
39. S. Borsanyi *et al.*, *JHEP* **1011**, 077 (2010).
40. G. Mangano *et al.*, *Phys. Lett.* **B534**, 8 (2002).
41. A. Linde, *Phys. Rev.* **D14**, 3345 (1976).
42. A. Linde, *Rept. on Prog. in Phys.* **42**, 389 (1979).
43. C.E. Vayonakis, *Surv. High Energy Physics* **5**, 87 (1986).
44. S.A. Bonometto and A. Masiero, *Nuovo Cimento* **9N5**, 1 (1986).
45. A. Linde, *Particle Physics And Inflationary Cosmology*, Harwood (1990).
46. K.A. Olive, *Phys. Reports* **190**, 307 (1990).
47. D. Lyth and A. Riotto, *Phys. Reports* **314**, 1 (1999).
48. A.H. Guth, *Phys. Rev.* **D23**, 347 (1981).
49. A.D. Linde, *Phys. Lett.* **108B**, 389 (1982).
50. A. Albrecht and P.J. Steinhardt, *Phys. Rev. Lett.* **48**, 1220 (1982).
51. S. Weinberg, *Rev. Mod. Phys.* **60**, 1 (1989).
52. L. Susskind, *hep-th/0302219* (2003).
53. B. Carr, *Universe or multiverse?* C.U.P. (2007).
54. A.D. Sakharov, *Sov. Phys. JETP Lett.* **5**, 24 (1967).
55. S. Weinberg, *Phys. Rev. Lett.* **42**, 850 (1979).

56. D. Toussaint *et al.*, Phys. Rev. **D19**, 1036 (1979).
57. I. Affleck and M. Dine, Nucl. Phys. **B249**, 361 (1985).
58. V. Kuzmin *et al.*, Phys. Lett. **B155**, 36 (1985).
59. M. Fukugita and T. Yanagida, Phys. Lett. **B174**, 45 (1986).
60. R. Cooke *et al.*, Astrophys. J. **781**, 31 (2013).
61. R.H. Cyburt *et al.*, Phys. Lett. **B567**, 227 (2003).
62. R.H. Cyburt *et al.*, JCAP **0811**, 012 (2008).
63. K.A. Olive *et al.*, Phys. Reports **333**, 389 (2000).
64. F. Iocco *et al.*, Phys. Reports **472**, 1 (2009).
65. R.H. Cyburt *et al.*, [arXiv:1505.01076](#).
66. D.J. Fixsen *et al.*, Astrophys. J. **473**, 576 (1996).
67. J.C. Mather *et al.*, Astrophys. J. **512**, 511 (1999).
68. S.M. Cole *et al.*, MNRAS **326**, 255 (2001).
69. A. Mesinger and Z. Haiman, Astrophys. J. **660**, 923 (2007).
70. S.D.M. White *et al.*, Nature **366**, 429 (1993).
71. S.W. Allen *et al.*, MNRAS **334**, L11 (2002).
72. S.W. Allen, MNRAS **296**, 392 (1998).
73. M. Kilbinger *et al.*, MNRAS **430**, 2200 (2013).
74. Planck Collab. 2015 Results XV, [arXiv:1502.01591](#).
75. S.M. Cole *et al.*, MNRAS **362**, 505 (2005).
76. W.J. Percival *et al.*, Astrophys. J. **657**, 645 (2007).
77. L. Anderson *et al.*, MNRAS **427**, 3435 (2012).
78. E. Linder, Phys. Rev. **D72**, 43529 (2005).
79. G. Efstathiou and J.R. Bond, MNRAS **218**, 103 (1986).
80. C. Gordon and A. Lewis, Phys. Rev. **D67**, 123513 (2003).
81. A. Vikhlinin *et al.*, Astrophys. J. **692**, 1060 (2009).
82. A. Dekel and O. Lahav, Astrophys. J. **520**, 24 (1999).
83. W.J. Percival *et al.*, MNRAS **381**, 1053 (2007).
84. W.J. Percival *et al.*, MNRAS **401**, 2148 (2010).
85. G. Efstathiou and J.R. Bond, MNRAS **304**, 75 (1999).
86. G.P. Efstathiou *et al.*, MNRAS **330**, L29 (2002).
87. Planck Collab. 2015 Results XX, [arXiv:1502.02114](#).
88. A. A. Starobinsky, Phys. Lett. **B91**, 99 (1980).
89. F. Bezrukov and M. Shaposhnikov, JHEP **0907**, 089 (2009).
90. Planck Collab. 2015 Results XXI, [arXiv:1502.01595](#).
91. N. Aghanim *et al.*, [arXiv:1502.01596](#).
92. W. Hu and I. Sawicki, Phys. Rev. **D76**, 4043 (2007).
93. B. Jain and P. Zhang, Phys. Rev. **D78**, 3503 (2008).
94. D.L. Wiltshire, Phys. Rev. Lett. **99**, 251101 (2007).
95. T. Buchert, Gen. Rel. Grav. **40**, 467 (2008).
96. I. Zlatev *et al.*, Phys. Rev. Lett. **82**, 896 (1999).
97. C. Armendariz-Picon, V. Mukhanov, and P.J. Steinhardt, Phys. Rev. **D63**, 3510 (2001).
98. S. DeDeo, R.R. Caldwell, and P.J. Steinhardt, Phys. Rev. **D67**, 3509 (2003).
99. W. Hu, [arXiv:astro-ph/0407158](#) (2004).
100. S.F. Daniel *et al.*, Phys. Rev. **D81**, 123508 (2010).

## 23. INFLATION

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### 23.1. Motivation and Introduction

The standard Big-Bang model of cosmology provides a successful framework in which to understand the thermal history of our Universe and the growth of cosmic structure, but it is essentially incomplete. As described in Sec. 22.2.4 in “Big Bang Cosmology” review, Big-Bang cosmology requires very specific initial conditions. It postulates a uniform cosmological background, described by a spatially-flat, homogeneous and isotropic Robertson-Walker (RW) metric (Eq. (22.1) in “Big Bang Cosmology” review), with scale factor  $R(t)$ . Within this setting, it also requires an initial almost scale-invariant distribution of primordial density perturbations as seen, for example, in the cosmic microwave background (CMB) radiation (described in Chap. 28, “Cosmic Microwave Background” review), on scales far larger than the causal horizon at the time the CMB photons last scattered.

The Hubble expansion rate,  $H \equiv \dot{R}/R$ , in a Robertson-Walker cosmology is given by the Friedmann constraint equation (Eq. (22.8) in “Big Bang Cosmology” review)

$$H^2 = \frac{8\pi\rho}{3M_P^2} + \frac{\Lambda}{3} - \frac{k}{R^2}, \quad (23.1)$$

where  $k/R^2$  is the intrinsic spatial curvature. We use natural units such that the speed of light  $c = 1$  and hence we have the Planck mass  $M_P = G_N^{-1/2} \simeq 10^{19}$  GeV. A cosmological constant,  $\Lambda$ , of the magnitude required to accelerate the Universe today (see Chap. 27, “Dark Energy” review) would have been completely negligible in the early Universe where the energy density  $\rho \gg M_P^2 \Lambda \sim 10^{-12}(\text{eV})^4$ . The standard early Universe cosmology, described in Sec. 22.1.5 in “Big Bang Cosmology” review, is thus dominated by non-relativistic matter ( $p_m = 0$ ) or radiation ( $p_r = \rho_r/3$  for an isotropic distribution). This leads to a decelerating expansion with  $\ddot{R} < 0$ .

The hypothesis of inflation [1,2] postulates a period of accelerated expansion,  $\ddot{R} > 0$ , in the very early Universe, preceding the standard radiation-dominated era, which offers a physical model for the origin of these initial conditions, as reviewed in [3,4,5,6,7]. Such a period of accelerated expansion (i) drives a curved RW spacetime (with spherical or hyperbolic spatial geometry) towards spatial flatness, and (ii) it also expands the causal horizon beyond the present Hubble length, so as to encompass all the scales relevant to describe the large-scale structure observed in our Universe today, via the following two mechanisms.

- (i) A spatially-flat universe with vanishing spatial curvature,  $k = 0$ , has the dimensionless density parameter  $\Omega_{\text{tot}} = 1$ , where we define (Eq. (22.13) in “Big Bang Cosmology” review; see Chap. 25, “Cosmological Parameters” review for more complete definitions)

$$\Omega_{\text{tot}} \equiv \frac{8\pi\rho_{\text{tot}}}{3M_P^2 H^2}, \quad (23.2)$$

with  $\rho_{\text{tot}} \equiv \rho + \Lambda M_P^2/8\pi$ . If we re-write the Friedmann constraint (Eq. (23.1)) in terms of  $\Omega_{\text{tot}}$  we have

$$1 - \Omega_{\text{tot}} = -\frac{k}{R^2}. \quad (23.3)$$

Observations require  $|1 - \Omega_{\text{tot},0}| < 0.005$  today [8], where the subscript 0 denotes the present-day value. Taking the time derivative of Eq. (23.3) we obtain

$$\frac{d}{dt}(1 - \Omega_{\text{tot}}) = -2\frac{\dot{R}}{R}(1 - \Omega_{\text{tot}}). \quad (23.4)$$

Thus in a decelerating expansion,  $\dot{R} > 0$  and  $\ddot{R} < 0$ , any small initial deviation from spatial flatness grows,  $(d/dt)|1 - \Omega_{\text{tot}}| > 0$ . A small value such as  $|1 - \Omega_{\text{tot},0}| < 0.005$  today requires an even smaller value at earlier times, e.g.,  $|1 - \Omega_{\text{tot}}| < 10^{-5}$  at the last scattering of the CMB, which appears unlikely, unless for

some reason space is exactly flat. However, an extended period of accelerated expansion in the very early Universe, with  $\dot{R} > 0$  and  $\ddot{R} > 0$  and hence  $(d/dt)|1 - \Omega_{\text{tot}}| < 0$ , can drive  $\Omega_{\text{tot}}$  sufficiently close to unity, so that  $|1 - \Omega_{\text{tot},0}|$  remains unobservably small today, even after the radiation- and matter-dominated eras, for a wide range of initial values of  $\Omega_{\text{tot}}$ .

- (ii) The comoving distance (the present-day proper distance) traversed by light between cosmic time  $t_1$  and  $t_2$  in an expanding universe can be written, (see Eq. (22.31) in “Big Bang Cosmology” review), as

$$D_0(t_1, t_2) = R_0 \int_{t_1}^{t_2} \frac{dt}{R(t)} = R_0 \int_{\ln R_1}^{\ln R_2} \frac{d(\ln R)}{\dot{R}}. \quad (23.5)$$

In standard decelerated (radiation- or matter-dominated) cosmology the integrand,  $1/\dot{R}$ , decreases towards the past, and there is a finite comoving distance traversed by light (a particle horizon) since the Big Bang ( $R_1 \rightarrow 0$ ). For example, the comoving size of the particle horizon at the CMB last-scattering surface ( $R_2 = R_{\text{ls}}$ ) corresponds to  $D_0 \sim 100$  Mpc, or approximately  $1^\circ$  on the CMB sky today (see Sec. 22.2.4 in “Big Bang Cosmology” review).

However, during a period of inflation,  $1/\dot{R}$  increases towards the past, and hence the integral (Eq. (23.5)) diverges as  $R_1 \rightarrow 0$ , allowing an arbitrarily large causal horizon, dependent only upon the duration of the accelerated expansion. Assuming that the Universe inflates with a finite Hubble rate  $H_*$  at  $t_1 = t_*$ , ending with  $H_{\text{end}} < H_*$  at  $t_2 = t_{\text{end}}$ , we have

$$D_0(t_*, t_{\text{end}}) > \left(\frac{R_0}{R_{\text{end}}}\right) H_*^{-1} (e^{N_*} - 1), \quad (23.6)$$

where  $N_* \equiv \ln(R_{\text{end}}/R_*)$  describes the duration of inflation, measured in terms of the logarithmic expansion (or “e-folds”) from  $t_1 = t_*$  up to the end of inflation at  $t_2 = t_{\text{end}}$ , and  $R_0/R_{\text{end}}$  is the subsequent expansion from the end of inflation to the present day. If inflation occurs above the TeV scale, the comoving Hubble scale at the end of inflation,  $(R_0/R_{\text{end}})H_{\text{end}}^{-1}$ , is less than one astronomical unit ( $\sim 10^{11}$  m), and a causally-connected patch can encompass our entire observable Universe today, which has a size  $D_0 > 30$  Gpc, if there were more than 40 e-folds of inflation ( $N_* > 40$ ). If inflation occurs at the GUT scale ( $10^{15}$  GeV) then we require more than 60 e-folds.

Producing an accelerated expansion in general relativity requires an energy-momentum tensor with negative pressure,  $p < -\rho/3$  (see Eq. (22.9) in “Big Bang Cosmology” review and Chap. 27, “Dark Energy” review), quite different from the hot dense plasma of relativistic particles in the hot Big Bang. However a positive vacuum energy  $V > 0$  does exert a negative pressure,  $p_V = -\rho_V$ . The work done by the cosmological expansion must be negative in this case so that the local vacuum energy density remains constant in an expanding universe,  $\dot{\rho}_V = -3H(\rho_V + p_V) = 0$ . Therefore, a false vacuum state can drive an exponential expansion, corresponding to a de Sitter spacetime with a constant Hubble rate  $H^2 = 8\pi\rho_V/3M_P^2$  on spatially-flat hypersurfaces.

A constant vacuum energy  $V$ , equivalent to a cosmological constant  $\Lambda$  in the Friedmann equation, cannot provide a complete description of inflation in the early Universe, since inflation must necessarily have come to an end in order for the standard Big-Bang cosmology to follow. A phase transition to the present true vacuum is required to release the false vacuum energy into the energetic plasma of the hot Big Bang and produce the large total entropy of our observed Universe today. Thus we must necessarily study dynamical models of inflation, where the time-invariance of the false vacuum state is broken by a time-dependent field. A first-order phase transition would produce a very inhomogeneous Universe [9] unless a time-dependent scalar field leads to a rapidly changing percolation rate [10,11,12]. However, a second-order phase transition [13,14], controlled by a slowly-rolling scalar field, can lead to a smooth classical exit from the vacuum-dominated phase.

As a spectacular bonus, quantum fluctuations in that scalar field could provide a source of almost scale-invariant density

fluctuations [15,16], as detected in the CMB (see section CosmicMicrowaveBackground), which are thought to be the origin of the structures seen in the Universe today.

Accelerated expansion and primordial perturbations can also be produced in some modified gravity theories (e.g., [1,17]), which introduce additional non-minimally coupled degrees of freedom. Such inflation models can often be conveniently studied by transforming variables to an ‘Einstein frame’ in which Einstein’s equations apply with minimally coupled scalar fields [18,19,20].

In the following we will review scalar field cosmology in general relativity and the spectra of primordial fluctuations produced during inflation, before studying selected inflation models.

## 23.2. Scalar Field Cosmology

The energy-momentum tensor for a canonical scalar field  $\phi$  with self-interaction potential  $V(\phi)$  is given in Eq. (22.51) in “Big Bang Cosmology” review. In a homogeneous background this corresponds to a perfect fluid with density

$$\rho = \frac{1}{2}\dot{\phi}^2 + V(\phi), \quad (23.7)$$

and isotropic pressure

$$p = \frac{1}{2}\dot{\phi}^2 - V(\phi), \quad (23.8)$$

while the 4-velocity is proportional to the gradient of the field,  $u^\mu \propto \nabla^\mu \phi$ .

A field with vanishing potential energy acts like a stiff fluid with  $p = \rho = \dot{\phi}^2/2$ , whereas if the time-dependence vanishes we have  $p = -\rho = -V$  and the scalar field is uniform in time and space. Thus a classical, potential-dominated scalar-field cosmology, with  $p \simeq -\rho$ , can naturally drive a quasi-de Sitter expansion; the slow time-evolution of the energy density weakly breaks the exact  $O(1,3)$  symmetry of four-dimensional de Sitter spacetime down to a Robertson-Walker (RW) spacetime, where the scalar field plays the role of the cosmic time coordinate.

In a scalar-field RW cosmology the Friedmann constraint equation (Eq. (23.1)) reduces to

$$H^2 = \frac{8\pi}{3M_P^2} \left( \frac{1}{2}\dot{\phi}^2 + V \right) - \frac{k}{R^2}, \quad (23.9)$$

while energy conservation (Eq. (22.10) in “Big Bang Cosmology” review) for a homogeneous scalar field reduces to the Klein-Gordon equation of motion (Eq. (22.53) in “Big Bang Cosmology” review)

$$\ddot{\phi} = -3H\dot{\phi} - V'(\phi). \quad (23.10)$$

The evolution of the scalar field is thus driven by the potential gradient  $V' = dV/d\phi$ , subject to damping by the Hubble expansion  $H\dot{\phi}$ .

If we define the Hubble slow-roll parameter

$$\epsilon_H \equiv -\frac{\dot{H}}{H^2}, \quad (23.11)$$

then we see that inflation ( $\ddot{R} > 0$  and hence  $\dot{H} > -H^2$ ) requires  $\epsilon_H < 1$ . In this case the spatial curvature decreases relative to the scalar field energy density as the Universe expands. Hence in the following we drop the spatial curvature and consider a spatially-flat RW cosmology, assuming that inflation has lasted sufficiently long that our observable universe is very close to spatially flatness. However, we note that bubble nucleation, leading to a first-order phase transition during inflation, can lead to homogeneous hypersurfaces with a hyperbolic (‘open’) geometry, effectively resetting the spatial curvature inside the bubble [21]. This is the basis of so-called open inflation models [22,23,24], where inflation inside the bubble has a finite duration, leaving a finite negative spatial curvature.

In a scalar field-dominated cosmology (Eq. (23.11)) gives

$$\epsilon_H = \frac{3\dot{\phi}^2}{2V + \dot{\phi}^2}, \quad (23.12)$$

in which case we see that inflation requires a potential-dominated expansion,  $\dot{\phi}^2 < V$ .

### 23.2.1. Slow-Roll Inflation :

It is commonly assumed that the field acceleration term,  $\ddot{\phi}$ , in (Eq. (23.10)) can be neglected, in which case one can give an approximate solution for the inflationary attractor [25]. This slow-roll approximation reduces the second-order Klein-Gordon equation (Eq. (23.10)) to a first-order system, which is over-damped, with the potential gradient being approximately balanced against the Hubble damping:

$$3H\dot{\phi} \simeq -V', \quad (23.13)$$

and at the same time that the Hubble expansion (Eq. (23.9)) is dominated by the potential energy

$$H^2 \simeq \frac{8\pi}{3M_P^2} V(\phi), \quad (23.14)$$

corresponding to  $\epsilon_H \ll 1$ .

A necessary condition for the validity of the slow-roll approximation is that the potential slow-roll parameters

$$\epsilon \equiv \frac{M_P^2}{16\pi} \left( \frac{V'}{V} \right)^2, \quad \eta \equiv \frac{M_P^2}{8\pi} \left( \frac{V''}{V} \right), \quad (23.15)$$

are small, i.e.,  $\epsilon \ll 1$  and  $|\eta| \ll 1$ , requiring the potential to be correspondingly flat. If we identify  $V''$  with the effective mass of the field, we see that the slow-roll approximation requires that the mass of the scalar field must be small compared with the Hubble scale. We note that the Hubble slow-roll parameter coincides with the potential slow-roll parameter,  $\epsilon_H \simeq \epsilon$ , to leading order in the slow-roll approximation.

The slow-roll approximation allows one to determine the Hubble expansion rate as a function of the scalar field value, and vice versa. In particular, we can express, in terms of the scalar field value during inflation, the total logarithmic expansion, or number of “e-folds”:

$$N_* \equiv \ln \left( \frac{R_{\text{end}}}{R_*} \right) = \int_{t_*}^{t_{\text{end}}} H dt \simeq - \int_{\phi_*}^{\phi_{\text{end}}} \sqrt{\frac{4\pi}{\epsilon}} \frac{d\phi}{M_P} \text{ for } V' > 0. \quad (23.16)$$

Given that the slow-roll parameters are approximately constant during slow-roll inflation,  $d\epsilon/dN \simeq 2\epsilon(\eta - 2\epsilon) = \mathcal{O}(\epsilon^2)$ , we have

$$N_* \simeq \frac{4}{\sqrt{\epsilon}} \frac{\Delta\phi}{M_P}. \quad (23.17)$$

Since we require  $N > 40$  to solve the flatness, horizon and entropy problems of the standard Big Bang cosmology, we require either very slow roll,  $\epsilon < 0.01$ , or a large change in the value of the scalar field relative to the Planck scale,  $\Delta\phi > M_P$ .

### 23.2.2. Reheating :

Slow-roll inflation can lead to an exponentially large universe, close to spatial flatness and homogeneity, but the energy density is locked in the potential energy of the scalar field, and needs to be converted to particles and thermalised to recover a hot Big Bang cosmology at the end of inflation [26,27]. This process is usually referred to as reheating, although there was not necessarily any preceding thermal era. Reheating can occur when the scalar field evolves towards the minimum of its potential, converting the potential energy first to kinetic energy. This can occur either through the breakdown of the slow-roll condition in single-field models, or due to an instability triggered by the inflaton reaching a critical value, in multi-field models known as hybrid inflation models [28].

Close to a simple minimum, the scalar field potential can be described by a quadratic function,  $V = m^2\phi^2/2$ , where  $m$  is the mass of the field. We can obtain slow-roll inflation in such a potential at large field values,  $\phi \gg M_P$ . However, for  $\phi \ll M_P$  the field approaches an oscillatory solution:

$$\phi(t) \simeq \frac{M_P}{\sqrt{3\pi}} \frac{\sin(mt)}{mt}. \quad (23.18)$$

For  $|\phi| < M_P$  the Hubble rate drops below the inflaton mass,  $H < m$ , and the field oscillates many times over a Hubble time. Averaging



over several oscillations,  $\Delta t \gg m^{-1}$ , we find  $\langle \dot{\phi}^2/2 \rangle_{\Delta t} \simeq \langle m^2 \phi^2/2 \rangle_{\Delta t}$  and hence

$$\langle \rho \rangle_{\Delta t} \simeq \frac{M_P^2}{6\pi t^2}, \quad \langle p \rangle_{\Delta t} \simeq 0. \quad (23.19)$$

This coherent oscillating field corresponds to a condensate of non-relativistic massive inflaton particles, driving a matter-dominated era at the end of inflation, with scale factor  $R \propto t^{2/3}$ .

The inflaton condensate can lose energy through perturbative decays due to terms in the interaction Lagrangian, such as

$$\mathcal{L}_{\text{int}} \subset -\lambda_i \sigma \phi \chi_i^2 - \lambda_j \phi \bar{\psi}_j \psi_j \quad (23.20)$$

that couple the inflation to scalar fields  $\chi_i$  or fermions  $\psi_j$ , where  $\sigma$  has dimensions of mass and the  $\lambda_i$  are dimensionless couplings. When the mass of the inflaton is much larger than the decay products, the decay rate is given by [29]

$$\Gamma_i = \frac{\lambda_i^2 \sigma^2}{8\pi m}, \quad \Gamma_j = \frac{\lambda_j^2 m}{8\pi}. \quad (23.21)$$

These decay products must in turn thermalise with Standard Model particles before we recover conventional hot Big Bang cosmology. An upper limit on the reheating temperature after inflation is given by [27]

$$T_{rh} = 0.2 \left( \frac{100}{g_*} \right)^{1/4} \sqrt{M_P \Gamma_{\text{tot}}}, \quad (23.22)$$

where  $g_*$  is the effective number of degrees of freedom and  $\Gamma_{\text{tot}}$  is the total decay rate for the inflaton, which is required to be less than  $m$  for perturbative decay.

The baryon asymmetry of the Universe must be generated after the main release of entropy during inflation, which is an important constraint on possible models. Also, the fact that the inflaton mass is much larger than the mass scale of the Standard Model opens up the possibility that it may decay into massive stable or metastable particles that could be connected with dark matter, constraining possible models. For example, in the context of supergravity models the reheat temperature is constrained by the requirement that gravitinos are not overproduced, potentially destroying the successes of Big Bang nucleosynthesis. For a range of gravitino masses one must require  $T_{rh} < 10^9$  GeV [30,31].

The process of inflaton decay and reheating can be significantly altered by interactions leading to space-time dependences in the effective masses of the fields. In particular, parametric resonance can lead to explosive, non-perturbative decay of the inflaton in some cases, a process often referred to as preheating [32,26]. For example, an interaction term of the form

$$\mathcal{L}_{\text{int}} \subset -\lambda^2 \phi^2 \chi^2, \quad (23.23)$$

leads to a time-dependent effective mass for the  $\chi$  field as the inflaton  $\phi$  oscillates. This can lead to non-adiabatic particle production if the bare mass of the  $\chi$  field is small for large couplings or for rapid changes of the inflaton field. The process of preheating is highly model-dependent, but it highlights the possible role of non-thermal particle production after and even during inflation.

### 23.3. Primordial Perturbations from Inflation

Although inflation was originally discussed as a solution to the problem of initial conditions required for homogeneous and isotropic hot Big Bang cosmology, it was soon realised that inflation also offered a mechanism to generate the inhomogeneous initial conditions required for the formation of large-scale structure [15,16,17,33].

#### 23.3.1. Metric Perturbations :

In a homogeneous classical inflationary cosmology driven by a scalar field, the inflaton field is uniform on constant-time hypersurfaces,  $\phi = \phi_0(t)$ . However, quantum fluctuations inevitably break the spatial symmetry leading to an inhomogeneous field:

$$\phi(t, x^i) = \phi_0(t) + \delta\phi(t, x^i). \quad (23.24)$$

At the same time, one should consider inhomogeneous perturbations of the RW spacetime metric (see, e.g., [34,35,36]) :

$$ds^2 = (1+2A)dt^2 - 2RB_i dt dx^i - R^2 [(1+2C)\delta_{ij} + \partial_i \partial_j E + h_{ij}] dx^i dx^j, \quad (23.25)$$

where  $A$ ,  $B$ ,  $E$  and  $C$  are scalar perturbations while  $h_{ij}$  represents transverse and tracefree, tensor metric perturbations. Vector metric perturbations can be eliminated using Einstein constraint equations in a scalar field cosmology.

The tensor perturbations remain invariant under a temporal gauge transformation  $t \rightarrow t + \delta t(t, x^i)$ , but both the scalar field and the scalar metric perturbations transform. For example, we have

$$\delta\phi \rightarrow \delta\phi - \dot{\phi}_0 \delta t, \quad C \rightarrow C - H \delta t. \quad (23.26)$$

However, there are gauge invariant combinations, such as [37]

$$Q = \delta\phi - \frac{\dot{\phi}_0}{H} C, \quad (23.27)$$

which describes the scalar field perturbations on spatially-flat hypersurfaces. This is simply related to the curvature perturbation on uniform-field hypersurfaces:

$$\mathcal{R} = C - \frac{H}{\dot{\phi}_0} \delta\phi = -\frac{H}{\dot{\phi}_0} Q, \quad (23.28)$$

which coincides in slow-roll inflation,  $\rho \simeq \rho(\phi)$ , with the curvature perturbation on uniform-density hypersurfaces [16]

$$\zeta = C - \frac{H}{\dot{\rho}_0} \delta\rho. \quad (23.29)$$

Thus scalar field and scalar metric perturbations are coupled by the evolution of the inflaton field.

#### 23.3.2. Gravitational waves from inflation :

The tensor metric perturbation,  $h_{ij}$  in Eq. (Eq. (23.25)), is gauge-invariant and decoupled from the scalar perturbations at first order. This represents the free excitations of the spacetime, i.e., gravitational waves, which are the simplest metric perturbations to study at linear order.

Each tensor mode, with wavevector  $\vec{k}$ , has two linearly-independent transverse and trace-free polarisation states:

$$h_{ij}(\vec{k}) = h_{\vec{k}} q_{ij} + \bar{h}_{\vec{k}} \bar{q}_{ij}. \quad (23.30)$$

The linearised Einstein equations then yield the same evolution equation for the amplitude as that for a massless field in RW spacetime:

$$\ddot{h}_{\vec{k}} + 3H\dot{h}_{\vec{k}} + \frac{k^2}{R^2} h_{\vec{k}} = 0, \quad (23.31)$$

(and similarly for  $\bar{h}_{\vec{k}}$ ). This can be re-written in terms of the conformal time,  $\eta = \int dt/R$ , and the conformally rescaled field:

$$u_{\vec{k}} = \frac{M_P R h_{\vec{k}}}{\sqrt{32\pi}}. \quad (23.32)$$

This conformal field then obeys the wave equation for a canonical scalar field in Minkowski spacetime with a time-dependent mass:

$$u_{\vec{k}}'' + \left( k^2 - \frac{R''}{R} \right) u_{\vec{k}} = 0. \quad (23.33)$$

During slow-roll

$$\frac{R''}{R} \simeq (2 - \epsilon) R^2 H^2. \quad (23.34)$$

This makes it possible on sub-Hubble scales,  $k^2/R^2 \gg H^2$ , where the background expansion can be neglected, to quantise the linearised metric fluctuations,  $u_{\vec{k}} \rightarrow \hat{u}_{\vec{k}}$ .

Crucially, in an inflationary expansion, where  $\ddot{R} > 0$ , the comoving Hubble length  $H^{-1}/R = 1/\dot{R}$  decreases with time. Thus all modes

start inside the Hubble horizon and it is possible to take the initial field fluctuations to be in a vacuum state at early times or on small scales:

$$\langle u_{\vec{k}_1} u_{\vec{k}_2} \rangle = \frac{i}{2} (2\pi)^3 \delta^{(3)}(\vec{k}_1 + \vec{k}_2). \quad (23.35)$$

In terms of the amplitude of the tensor metric perturbations, this corresponds to

$$\langle h_{\vec{k}_1} h_{\vec{k}_2} \rangle = \frac{1}{2} \frac{\mathcal{P}_T(k_1)}{4\pi k_1^3} (2\pi)^3 \delta^{(3)}(\vec{k}_1 + \vec{k}_2), \quad (23.36)$$

where the factor 1/2 appears due to the two polarisation states that contribute to the total tensor power spectrum:

$$\mathcal{P}_t(k) = \frac{64\pi}{M_P^2} \left( \frac{k}{2\pi R} \right)^2. \quad (23.37)$$

On super-Hubble scales,  $k^2/R^2 \ll H^2$ , we have the growing mode solution,  $u_{\vec{k}} \propto R$ , corresponding to  $h_{\vec{k}} \rightarrow \text{constant}$ , i.e., tensor modes are frozen-in on super-Hubble scales, both during and after inflation. Thus, connecting the initial vacuum fluctuations on sub-Hubble scales to the late-time power spectrum for tensor modes at Hubble exit during inflation,  $k = R_* H_*$ , we obtain

$$\mathcal{P}_t(k) \simeq \frac{64\pi}{M_P^2} \left( \frac{H_*}{2\pi} \right)^2. \quad (23.38)$$

In the de Sitter limit,  $\epsilon \rightarrow 0$ , the Hubble rate becomes time-independent and the tensor spectrum on super-Hubble scales becomes scale-invariant [38]. However slow-roll evolution leads to weak time dependence of  $H_*$  and thus a scale-dependent spectrum on large scales, with a spectral tilt

$$n_t \equiv \frac{d \ln \mathcal{P}_T}{d \ln k} \simeq -2\epsilon_*. \quad (23.39)$$

### 23.3.3. Density Perturbations from Inflation :

The scalar field fluctuations on spatially-flat hypersurfaces are coupled to scalar metric perturbations at first order, but these can be eliminated using the Einstein constraint equations to yield an evolution equation

$$\ddot{Q}_{\vec{k}} + 3H\dot{Q}_{\vec{k}} + \left[ \frac{k^2}{R^2} + V'' - \frac{8\pi}{3M_P^2} \frac{d}{dt} \left( \frac{R^3 \dot{\phi}^2}{H} \right) \right] Q_{\vec{k}} = 0. \quad (23.40)$$

Terms proportional to  $M_P^{-2}$  represent the effect on the field fluctuations of gravity at first order. As can be seen, this vanishes in the limit of a constant background field, and hence is suppressed in the slow-roll limit, but it is of the same order as the effective mass,  $V'' = 3\eta H^2$ , so must be included if we wish to model deviations from exact de Sitter symmetry.

This wave equation can also be written in the canonical form for a free field in Minkowski spacetime if we define [37]

$$v_{\vec{k}} \equiv R Q_{\vec{k}}, \quad (23.41)$$

to yield

$$v_{\vec{k}}'' + \left( k^2 - \frac{z''}{z} \right) v_{\vec{k}} = 0, \quad (23.42)$$

where we define

$$z \equiv \frac{R\dot{\phi}}{H}, \quad \frac{z''}{z} \simeq (2 + 5\epsilon - 3\eta) R^2 H^2, \quad (23.43)$$

where the last approximate equality holds to leading order in the slow-roll approximation.

As previously done for gravitational waves, we quantise the linearised field fluctuations  $v_{\vec{k}} \rightarrow \hat{v}_{\vec{k}}$  on sub-Hubble scales,  $k^2/R^2 \gg H^2$ , where the background expansion can be neglected. Thus we impose

$$\langle v_{\vec{k}_1} v_{\vec{k}_2} \rangle = \frac{i}{2} (2\pi)^3 \delta^{(3)}(\vec{k}_1 + \vec{k}_2). \quad (23.44)$$

In terms of the field perturbations, this corresponds to

$$\langle Q_{\vec{k}_1} Q_{\vec{k}_2} \rangle = \frac{\mathcal{P}_Q(k_1)}{4\pi k_1^3} (2\pi)^3 \delta^{(3)}(\vec{k}_1 + \vec{k}_2), \quad (23.45)$$

where the power spectrum for vacuum field fluctuations on sub-Hubble scales,  $k^2/R^2 \gg H^2$ , is simply

$$\mathcal{P}_Q(k) = \left( \frac{k}{2\pi R} \right)^2, \quad (23.46)$$

yielding the classic result for the vacuum fluctuations for a massless field in de Sitter at Hubble exit,  $k = R_* H_*$ :

$$\mathcal{P}_Q(k) \simeq \left( \frac{H}{2\pi} \right)_*^2. \quad (23.47)$$

In practice there are slow-roll corrections due to the small but finite mass ( $\eta$ ) and field evolution ( $\epsilon$ ) [39].

Slow-roll corrections to the field fluctuations are small on sub-Hubble scales, but can become significant as the field and its perturbations evolve over time on super-Hubble scales. Thus it is helpful to work instead with the curvature perturbation,  $\zeta$  defined in equation (Eq. (23.29)), which remains constant on super-Hubble scales for adiabatic density perturbations both during and after inflation [16,40]. Thus we have an expression for the primordial curvature perturbation on super-Hubble scales produced by single-field inflation:

$$\mathcal{P}_\zeta(k) = \left[ \left( \frac{H}{\dot{\phi}} \right)^2 \mathcal{P}_Q(k) \right]_* \simeq \frac{4\pi}{M_P^2} \left[ \frac{1}{\epsilon} \left( \frac{H}{2\pi} \right)^2 \right]_* \quad (23.48)$$

Comparing this with the primordial gravitational wave power spectrum (Eq. (23.38)) we obtain the tensor-to-scalar ratio for single-field slow-roll inflation

$$r \equiv \frac{\mathcal{P}_\zeta}{\mathcal{P}_t} \simeq 16\epsilon_*. \quad (23.49)$$

Note that the scalar amplitude is boosted by a factor  $1/\epsilon_*$  during slow-roll inflation, because small scalar field fluctuations can lead to relatively large curvature perturbations on hypersurfaces defined with respect to the density if the potential energy is only weakly dependent on the scalar field, as in slow-roll. Indeed, the de Sitter limit is singular, since the potential energy becomes independent of the scalar field at first order,  $\epsilon \rightarrow 0$ , and the curvature perturbation on uniform-density hypersurfaces becomes ill-defined.

We note that in single-field inflation the tensor-to-scalar ratio and the tensor tilt (Eq. (23.39)) at the same scale are both determined by the first slow-roll parameter at Hubble exit,  $\epsilon_*$ , giving rise to an important consistency test for single-field inflation:

$$n_t = -\frac{r}{8}. \quad (23.50)$$

This may be hard to verify if  $r$  is small, making any tensor tilt  $n_t$  difficult to measure. On the other hand, it does offer a way to rule out single-field slow-roll inflation if either  $r$  or  $n_t$  is large.

Given the relatively large scalar power spectrum, it has proved easier to measure the scalar tilt, conventionally defined as  $n_s - 1$ . Slow-roll corrections lead to slow time-dependence of both  $H_*$  and  $\epsilon_*$ , giving a weak scale-dependence of the scalar power spectrum:

$$n_s - 1 \equiv \frac{d \ln \mathcal{P}_\zeta}{d \ln k} \simeq -6\epsilon_* + 2\eta_*, \quad (23.51)$$

and a running of this tilt at second-order in slow-roll:

$$\frac{dn_s}{d \ln k} \simeq -8\epsilon_*(3\epsilon_* - 2\eta_*) - 2\epsilon_*^2, \quad (23.52)$$

where the running introduces a new slow-roll parameter at second-order:

$$\xi^2 \equiv \frac{M_P^4}{64\pi^2} \frac{V'V'''}{V^2}. \quad (23.53)$$

Any relation between the tensor-to-scalar ratio and the scalar tilt must impose some model-dependent relation between the slow-roll parameters. For example, for power-law inflation or chaotic inflation driven by a massive field (see later) we have  $\eta \simeq \epsilon$  and hence

$$n_s - 1 \simeq -\frac{r}{8}. \quad (23.54)$$

Violating this condition would rule out these specific classes of single-field models.

### 23.3.4. Observational Bounds :

The observed scale-dependence of the power spectrum makes it necessary to specify the comoving scale,  $k$ , at which quantities are constrained and hence the Hubble-exit time,  $k = a_* H_*$ , when the corresponding theoretical quantities are calculated during inflation. This is usually expressed in terms of the number of e-folds from the end of inflation [41]:

$$N_*(k) \simeq 67 - \ln\left(\frac{k}{a_0 H_0}\right) + \frac{1}{4} \ln\left(\frac{V_*^2}{M_{\text{Pl}}^4 \rho_{\text{end}}}\right) + \frac{1}{12} \ln\left(\frac{\rho_{rh}}{\rho_{\text{end}}}\right) - \frac{1}{12} \ln(g_*), \quad (23.55)$$

where  $H_0^{-1}/a_0$  is the present comoving Hubble length. Different models of reheating and thus different reheat temperatures and densities,  $\rho_{rh}$  in Eq. (23.55), lead to a range of possible values for  $N_*$  corresponding to a fixed physical scale, and hence we have a range of observational predictions for a given inflation model, as seen in Fig. 23.1.

The Planck 2015 temperature and polarisation data (see Chap. 28, “Cosmic Microwave Background” review) are consistent with a smooth featureless power spectrum over a range of comoving wavenumbers,  $0.008 h^{-1} \text{Mpc}^{-1} \leq k \leq 0.1 h \text{Mpc}^{-1}$ . In the absence of running, the data measure the spectral index

$$n_s = 0.968 \pm 0.006, \quad (23.56)$$

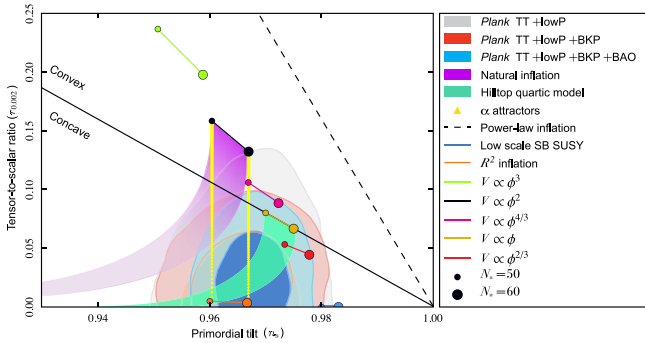
corresponding to a deviation from scale-invariance exceeding the  $5\sigma$  level. If running of the spectral tilt is included in the model, this is constrained to be

$$\frac{dn_s}{d \ln k} = -0.003 \pm 0.007. \quad (23.57)$$

A recent analysis of the BICEP2/Keck Array, Planck and other data places an upper bound on the tensor-to-scalar ratio [42]

$$r < 0.07 \quad (23.58)$$

at the 95% CL.



**Figure 23.1:** The marginalized joint 68 and 95% CL regions for the tilt in the scalar perturbation spectrum,  $n_s$ , and the relative magnitude of the tensor perturbations,  $r$ , obtained from the Planck 2015 data and their combinations with BICEP2/Keck Array and/or BAO data, confronted with the predictions of some of the inflationary models discussed in this review. This figure is taken from [44].

These observational bounds can be converted into bounds on the slow-roll parameters and hence the potential during slow-roll inflation. Setting higher-order slow-roll parameters (beyond second-order in horizon-flow parameters [43]) to zero the Planck collaboration obtain the following bounds [44]

$$\epsilon < 0.012, \quad (23.59)$$

$$\eta = -0.0080^{+0.0088}_{-0.0146}, \quad (23.60)$$

$$\xi^2 = 0.0070^{+0.0045}_{-0.0069}, \quad (23.61)$$

which can be used to constrain models, as discussed in the next Section.

Fig. 23.1, which is taken from [44], compares observational CMB constraints on the tilt,  $n_s$ , in the spectrum of scalar perturbations and the ratio,  $r$ , between the magnitudes of tensor and scalar perturbations. Important rôles are played by data from the Planck satellite, the BICEP2/Keck Array (BKP) and measurements of baryon acoustic oscillations (BAO). The reader is referred to [44] for technical details. These experimental constraints are compared with the predictions of some of the inflationary models discussed in this review. Generally speaking, models with a concave potential are favoured over those with a convex potential, and models with power-law inflation, as opposed to de Sitter-like (quasi-)exponential expansion, are now excluded.

## 23.4. Models

### 23.4.1. Pioneering Models :

The paradigm of the inflationary Universe was proposed in [2], where it was pointed out that an early period of (near-)exponential expansion, in addition to resolving the horizon and flatness problems of conventional Big-Bang cosmology as discussed above (the possibility of a de Sitter phase in the early history of the Universe was also proposed in the non-minimal gravity model of [1], with the motivation of avoiding an initial singularity), would also dilute the prior abundance of any unseen heavy, (meta-)stable particles, as exemplified by monopoles in grand unified theories (GUTs; see Chap. 16, “Grand Unified Theories” review). The original proposal was that this inflationary expansion took place while the Universe was in a metastable state (a similar suggestion was made in [45,46], where in [45] it was also pointed out that such a mechanism could address the horizon problem) and was terminated by a first-order transition due to tunnelling through a potential barrier. However, it was recognized already in [2] that this ‘old inflation’ scenario would need modification if the transition to the post-inflationary universe were to be completed smoothly without generating unacceptable inhomogeneities.

This ‘graceful exit’ problem was addressed in the ‘new inflation’ model of [13] (see also [14] and footnote [39] of [2]), which studied models based on an  $SU(5)$  GUT with an effective potential of the Coleman-Weinberg type (i.e., dominated by radiative corrections), in which inflation could occur during the roll-down from the local maximum of the potential towards a global minimum. However, it was realized that the Universe would evolve to a different minimum from the Standard Model [47], and it was also recognized that density fluctuations would necessarily be too large [15], since they were related to the GUT coupling strength.

These early models of inflation assumed initial conditions enforced by thermal equilibrium in the early Universe. However, this assumption was questionable: indeed, it was not made in the model of [1], in which a higher-order gravitational curvature term was assumed to arise from quantum corrections, and the assumption of initial thermal equilibrium was jettisoned in the ‘chaotic’ inflationary model of [48]. These are the inspirations for much recent inflationary model building, so we now discuss them in more detail, before reviewing contemporary models.

In this section we will work in natural units where we set the reduced Planck mass to unity, i.e.,  $8\pi/M_{\text{Pl}}^2 = 1$ . All masses are thus relative to the reduced Planck scale.

### 23.4.2. $R^2$ Inflation :

The first-order Einstein-Hilbert action,  $(1/2) \int d^4x \sqrt{-g} R$ , where  $R$  is the Ricci scalar curvature, is the minimal possible theory consistent with general coordinate invariance. However, it is possible that there might be non-minimal corrections to this action, and the unique second-order possibility is

$$S = \frac{1}{2} \int d^4x \sqrt{-g} \left( R + \frac{R^2}{6M^2} \right). \quad (23.62)$$

It was pointed out in [1] that an  $R^2$  term could be generated by quantum effects, and that (Eq. (23.62)) could lead to de Sitter-like

expansion of the Universe. Scalar density perturbations in this model were calculated in [17]. Because the initial phase was (almost) de Sitter, these perturbations were (approximately) scale-invariant, with magnitude  $\propto M$ . It was pointed out in [17] that requiring the scalar density perturbations to lie in the range  $10^{-3}$  to  $10^{-5}$ , consistent with upper limits at that time, would require  $M \sim 10^{-3}$  to  $10^{-5}$  in Planck units, and it was further suggested in that these perturbations could lead to the observed large-scale structure of the Universe, including the formation of galaxies.

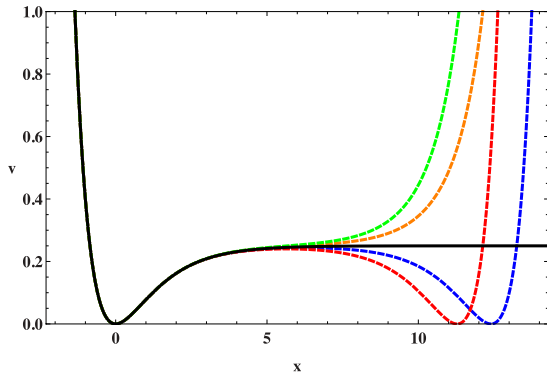
Although the action (Eq. (23.62)) does not contain an explicit scalar field, [17] reduced the calculation of density perturbations to that of fluctuations in the scalar curvature  $R$ , which could be identified (up to a factor) with a scalar field of mass  $M$ . The formal equivalence of  $R^2$  gravity (Eq. (23.62)) to a theory of gravity with a massive scalar  $\phi$  had been shown in [18], see also [19]. The effective scalar potential for what we would nowadays call the ‘inflaton’ [49] takes the form

$$S = \frac{1}{2} \int d^4x \sqrt{-g} \left[ R + (\partial_\mu \phi)^2 - \frac{3}{2} M^2 (1 - e^{-\sqrt{2/3} \phi})^2 \right] \quad (23.63)$$

when the action is written in the Einstein frame, and the potential is shown as the solid black line in Fig. 23.2. Using (Eq. (23.48)), one finds that the amplitude of the scalar density perturbations in this model is given by

$$\Delta \mathcal{R} = \frac{3M^2}{8\pi^2} \sinh^4 \left( \frac{\phi}{\sqrt{6}} \right), \quad (23.64)$$

The measured magnitude of the density fluctuations in the CMB requires  $M \simeq 1.3 \times 10^{-5}$  in Planck units (assuming  $N_* \simeq 55$ ), so one of the open questions in this model is why  $M$  is so small. Obtaining  $N_* \simeq 55$  also requires an initial value of  $\phi \simeq 5.5$ , i.e., a super-Planckian initial condition, and another issue for this and many other models is how the form of the effective potential is protected and remains valid at such large field values. Using Eq. (23.51) one finds that  $n_s \simeq 0.965$  for  $N_* \simeq 55$  and using (Eq. (23.49)) one finds that  $r \simeq 0.0035$ . These predictions are consistent with the present data from Planck and other experiments, as seen in Fig. 23.1.



**Figure 23.2:** The inflationary potential  $V$  in the  $R^2$  model (solid black line) compared with its form in various no-scale models discussed in detail in [50] (dashed coloured lines).

#### 23.4.3. Chaotic Models with Power-Law Potentials :

As has already been mentioned, a key innovation in inflationary model-building was the suggestion to abandon the questionable assumption of a thermal initial state, and consider ‘chaotic’ initial conditions with very general forms of potential [48]. (Indeed, the  $R^2$  model discussed above can be regarded as a prototype of this approach.) The chaotic approach was first proposed in the context of a simple power-law potential of the form  $\mu^{4-\alpha} \phi^\alpha$ , and the specific example of  $\lambda \phi^4$  was studied in [48]. Such models make the following predictions for the slow-roll parameters  $\epsilon$  and  $\eta$ :

$$\epsilon = \frac{1}{2} \left( \frac{\alpha}{\phi} \right)^2, \quad \eta = \frac{\alpha(\alpha-1)}{\phi^2}, \quad (23.65)$$

leading to the predictions

$$r \approx \frac{4\alpha}{N_*}, \quad n_s - 1 \approx -\frac{\alpha+2}{2N_*}, \quad (23.66)$$

which are shown in Fig. 23.1 for some illustrative values of  $\alpha$ . We note that the prediction of the original  $\phi^4$  model lies out of the frame, with values of  $r$  that are too large and values of  $n_s$  that are too small. The  $\phi^3$  model has similar problems, and would in any case require modification in order to have a well-defined minimum. The simplest possibility is  $\phi^2$ , but this is now also disfavoured by the data, at the 95% CL if only the Planck data are considered, and more strongly if other data are included, as seen in Fig. 23.1. (For non-minimal models of quadratic inflation that avoid this problem, see, e.g., [51].)

Indeed, as can be seen in Fig. 23.1, all models with a convex potential (i.e., one curving upwards) are disfavoured compared to models with a concave potential. Thus, a model with a  $\phi^{2/3}$  potential may just be compatible with the data at the 68% CL, whereas linear and  $\phi^{4/3}$  potentials are allowed only at about the 95% CL.

#### 23.4.4. Hilltop Models :

This preference for a concave potential motivates interest in ‘hilltop’ models [52], whose starting-point is a potential of the form

$$V(\phi) = \Lambda^4 \left[ 1 - \left( \frac{\phi}{\mu} \right)^p + \dots \right], \quad (23.67)$$

where the  $\dots$  represent extra terms that yield a positive semi-definite potential. To first order in the slow-roll parameters, when  $x \equiv \phi/\mu$  is small, one has

$$n_s \simeq 1 - p(p-1)\mu^{-2} \frac{x^{p-2}}{(1-x^p)} - \frac{3}{8}r, \quad r \simeq 8p^2\mu^{-2} \frac{x^{2p-2}}{(1-x^p)^2}. \quad (23.68)$$

As seen in Fig. 23.1, a hilltop model with  $p = 4$  can be compatible with the Planck and other measurements, if  $\mu \gg M_P$ .

#### 23.4.5. D-Brane Inflation :

Many scenarios for inflation involving extra dimensions have been proposed, e.g., the possibility that observable physics resides on a three-dimensional brane, and that there is an inflationary potential that depends on the distance between our brane and an antibrane, with a potential of the form [53]

$$V(\phi) = \Lambda^4 \left[ 1 - \left( \frac{\mu}{\phi} \right)^p + \dots \right]. \quad (23.69)$$

In this scenario the effective potential vanishes in the limit  $\phi \rightarrow \infty$ , corresponding to complete separation between our brane and the antibrane. The predictions for  $n_s$  and  $r$  in this model can be obtained from (Eq. (23.68)) by exchanging  $p \leftrightarrow -p$ , and are also consistent with the Planck and other data.

#### 23.4.6. Natural Inflation :

Also seen in Fig. 23.1 are the predictions of ‘natural inflation’ [54], in which one postulates a non-perturbative shift symmetry that suppresses quantum corrections, so that a hierarchically small scale of inflation,  $H \ll M_P$ , is technically natural. In the simplest models, there is a periodic potential of the form

$$V(\phi) = \Lambda^4 \left[ 1 + \cos \left( \frac{\phi}{f} \right) \right], \quad (23.70)$$

where  $f$  is a dimensional parameter reminiscent of an axion decay constant (see the next subsection) [55], which must have a value  $> M_P$ . Natural inflation can yield predictions similar to quadratic inflation (which are no longer favoured, as already discussed), but can also yield an effective convex potential. Thus, it may lead to values of  $r$  that are acceptably small, but for values of  $n_s$  that are in tension with the data, as seen in Fig. 23.1.

### 23.4.7. Axion Monodromy Models :

The effective potentials in stringy models [56,57] motivated by axion monodromy may be of the form

$$V(\phi) = \mu^{4-\alpha} \phi^\alpha + \Lambda^4 e^{-C \left( \frac{\phi}{\phi_0} \right)^{p_\Lambda}} \cos \left[ \gamma + \frac{\phi}{f} \left( \frac{\phi}{\phi_0} \right)^{p_f+1} \right], \quad (23.71)$$

where  $\mu, \Lambda, f$  and  $\phi_0$  are parameters with the dimension of mass, and  $C, p, p_\Lambda, p_f$  and  $\gamma$  are dimensionless constants, generalizing the potential ([54]) in the simplest models of natural inflation. The oscillations in (Eq. (23.71)) are associated with the axion field, and powers  $p_\Lambda, p_f \neq 0$  may arise from  $\phi$ -dependent evolutions of string moduli. Since the exponential prefactor in (Eq. (23.71)) is due to non-perturbative effects that may be strongly suppressed, the oscillations may be unobservably small. Specific string models having  $\phi^\alpha$  with  $\alpha = 4/3, 1$  or  $2/3$  have been constructed in [56,57], providing some motivation for the low-power models mentioned above.

As seen in Fig. 23.1, the simplest axion monodromy models with these values of the power  $\alpha$  are compatible with all the available data at the 95% CL, though not at the 68% CL. The Planck Collaboration has also searched for characteristic effects associated with the second term in (Eq. (23.71)), such as a possible drift in the modulation amplitude (setting  $p_\Lambda = C = 0$ ), and a possible drifting frequency generated by  $p_f \neq 0$ , without finding any compelling evidence [44].

### 23.4.8. Higgs Inflation :

Since the energy scale during inflation is commonly expected to lie between the Planck and TeV scales, it may serve as a useful bridge with contacts both to string theory or some other quantum theory of gravity, on the one side, and particle physics on the other side. However, as the above discussion shows, much of the activity in building models of inflation has been largely independent of specific connections with these subjects, though some examples of string-motivated models of inflation were mentioned above.

The most economical scenario for inflation might be to use as inflaton the only established scalar field, namely the Higgs field (see Chap. 11, “Status of Higgs boson physics” review). A specific model assuming a non-minimal coupling of the Higgs field  $h$  to gravity was constructed in [58]. Its starting-point is the action

$$S = \int d^4x \sqrt{-g} \left[ \frac{M^2 + \xi h^2}{2} R + \frac{1}{2} \partial_\mu h \partial^\mu h - \frac{\lambda}{4} (h^2 - v^2)^2 \right], \quad (23.72)$$

where  $v$  is the Higgs vacuum expectation value. The model requires  $\xi \gg 1$ , in which case it can be rewritten in the Einstein frame as

$$S = \int d^4x \sqrt{-g} \left[ \frac{1}{2} R + \frac{1}{2} \partial_\mu \chi \partial^\mu \chi - U(\chi) \right], \quad (23.73)$$

where the effective potential for the canonically-normalized inflaton field  $\chi$  has the form

$$U(\chi) = \frac{\lambda}{4\xi^2} \left[ 1 + \exp \left( -\frac{2\chi}{\sqrt{6}M_P} \right) \right]^{-2}, \quad (23.74)$$

which is similar to the effective potential of the  $R^2$  model at large field values. As such, the model inflates successfully if  $\xi \simeq 5 \times 10^4 m_h/(\sqrt{2}v)$ , with predictions for  $n_s$  and  $r$  that are indistinguishable from the predictions of the  $R^2$  model shown in Fig. 23.1.

This model is very appealing, but must confront several issues. One is to understand the value of  $\xi$ , and another is the possibility of unitarity violation. However, a more fundamental issue is whether the effective quartic Higgs coupling is positive at the scale of the Higgs field during inflation. Extrapolations of the effective potential in the Standard Model using the measured values of the masses of the Higgs boson and the top quark indicate that probably  $\lambda < 0$  at this scale [59], though there are still significant uncertainties associated with the appropriate input value of the top mass and the extrapolation to high renormalization scales.

### 23.4.9. Supersymmetric Models of Inflation :

Supersymmetry [60] is widely considered to be a well-motivated possible extension of the Standard Model that might become apparent at the TeV scale. It is therefore natural to consider supersymmetric models of inflation. These were originally proposed because of the problems of the the new inflationary theory [13,14] based on the one-loop (Coleman-Weinberg) potential for breaking SU(5). Several of these problems are related to the magnitude of the effective potential parameters: in any model of inflation based on an elementary scalar field, some parameter in the effective potential must be small in natural units, e.g., the quartic coupling  $\lambda$  in a chaotic model with a quartic potential, or the mass parameter  $\mu$  in a model of chaotic quadratic inflation. These parameters are renormalized multiplicatively in a supersymmetric theory, so that the quantum corrections to small values would be under control. Hence it was suggested that inflation cries out for supersymmetry [61], though non-supersymmetric resolutions of the problems of Coleman-Weinberg inflation are also possible: see, e.g., Ref. [62].

In the Standard Model there is only one scalar field that could be a candidate for the inflaton, namely the Higgs field discussed above, but even the minimal supersymmetric extension of the Standard Model (MSSM) contains many scalar fields. However, none of these is a promising candidate for the inflaton. The minimal extension of the MSSM that may contain a suitable candidate is the supersymmetric version of the minimal seesaw model of neutrino masses, which contains the three supersymmetric partners of the heavy singlet (right-handed) neutrinos. One of these singlet sneutrinos  $\tilde{\nu}$  could be the inflaton [63]: it would have a quadratic potential, the mass coefficient required would be  $\sim 10^{13}$  GeV, very much in the expected ball-park for singlet (right-handed) neutrino masses, and sneutrino inflaton decays also could give rise to the cosmological baryon asymmetry via leptogenesis. However, as seen in Fig. 23.1 and already discussed, a purely quadratic inflationary potential is no longer favoured by the data. This difficulty could in principle be resolved in models with multiple sneutrinos [64], or by postulating a trilinear sneutrino coupling and hence a superpotential of Wess-Zumino type [65], which can yield successful inflation with predictions intermediate between those of natural inflation and hilltop inflation in Fig. 23.1.

Finally, we note that it is also possible to obtain inflation via supersymmetry breaking, as in the model [66] whose predictions are illustrated in Fig. 23.1.

### 23.4.10. Supergravity Models :

Any model of early-Universe cosmology, and specifically inflation, must necessarily incorporate gravity. In the context of supersymmetry this requires an embedding in some supergravity theory [67,68]. An  $\mathcal{N} = 1$  supergravity theory is specified by three functions: a Hermitian function of the matter scalar fields  $\phi^i$ , called the Kähler potential  $K$ , that describes its geometry, a holomorphic function of the superfields, called the superpotential  $W$ , which describes their interactions, and another holomorphic function  $f_{\alpha\beta}$ , which describes their couplings to gauge fields  $V_\alpha$  [69].

The simplest possibility is that the Kähler metric is flat:

$$K = \phi^i \phi_i^*, \quad (23.75)$$

where the sum is over all scalar fields in the theory, and the simplest inflationary model in minimal supergravity had the superpotential [70]

$$W = m^2(1 - \phi)^2, \quad (23.76)$$

Where  $\phi$  is the inflaton. However, this model predicts a tilted scalar perturbation spectrum,  $n_s = 0.933$ , which is now in serious disagreement with the data from Planck and other experiments shown in Fig. 23.1.

Moreover, there is a general problem that arises in any supergravity theory coupled to matter, namely that, since its effective scalar potential contains a factor of  $e^K$ , scalars typically receive squared masses  $\propto H^2 \sim V$ , where  $H$  is the Hubble parameter [71], an issue called the ‘ $\eta$  problem’. The theory given by (Eq. (23.76)) avoids this  $\eta$  problem, but a generic supergravity inflationary model encounters

this problem of a large inflaton mass. Moreover, there are additional challenges for supergravity inflation associated with the spontaneous breaking of local supersymmetry [72,73,74].

Various approaches to the  $\eta$  problem in supergravity have been proposed, including the possibility of a shift symmetry [75], and one possibility that has attracted renewed attention recently is no-scale supergravity [76,77]. This is a form of supergravity with a Kähler potential that can be written in the form [78]

$$K = -3 \ln \left( T + T^* - \frac{\sum_i |\phi^i|^2}{3} \right), \quad (23.77)$$

which has the special property that it naturally has a flat potential, at the classical level and before specifying a non-trivial superpotential. As such, no-scale supergravity is well-suited for constructing models of inflation. Adding to its attraction is the feature that compactifications of string theory to supersymmetric four-dimensional models yield effective supergravity theories of the no-scale type [79]. There are many examples of superpotentials that yield effective inflationary potentials for either the  $T$  field (which is akin to a modulus field in some string compactification) or a  $\phi$  field (generically representing matter) that are of the same form as the effective potential of the  $R^2$  model (Eq. (23.63)) when the magnitude of the inflaton field  $\gg 1$  in Planck units, as required to obtain sufficiently many e-folds of inflation,  $N_*$  [80,81]. This framework also offers the possibility of using a suitable superpotential to construct models with effective potentials that are similar, but not identical, to the  $R^2$  model, as shown by the dashed coloured lines in Fig. 23.2.

#### 23.4.11. Other Exponential Potential Models :

This framework also offers the possibility [80] of constructing models in which the asymptotic constant value of the potential at large inflaton field values is approached via a different exponentially-suppressed term:

$$V(\phi) = A \left[ 1 - \delta e^{-B\phi} + \mathcal{O}(e^{-2B\phi}) \right], \quad (23.78)$$

where the magnitude of the scalar density perturbations fixes  $A$ , but  $\delta$  and  $B$  are regarded as free parameters. In the case of  $R^2$  inflation  $\delta = 2$  and  $B = \sqrt{2/3}$ . In a model such as (Eq. (23.78)), one finds at leading order in the small quantity  $e^{-B\phi}$  that

$$\begin{aligned} n_s &= 1 - 2B^2 \delta e^{-B\phi}, \\ r &= 8B^2 \delta^2 e^{-2B\phi}, \\ N_* &= \frac{1}{B^2 \delta} e^{+B\phi}. \end{aligned} \quad (23.79)$$

yielding the relations

$$n_s = 1 - \frac{2}{N_*}, r = \frac{8}{B^2 N_*^2}. \quad (23.80)$$

This model leads to the class of predictions labelled by ‘ $\alpha$  attractors’ [82] in Fig. 23.1. There are generalizations of the simplest no-scale model (Eq. (23.77)) with prefactors before the  $\ln(\dots)$  that are 1 or 2, leading to larger values of  $B = \sqrt{2}$  or 1, respectively, and hence smaller values of  $r$  than in the  $R^2$  model.

### 23.5. Model Comparison

Given a particular inflationary model, one can obtain constraints on the model parameters, informed by the likelihood, corresponding to the probability of the data given a particular choice of parameters (see Chap. 39, “Statistics” review). In the light of the detailed constraints on the statistical distribution of primordial perturbations now inferred from high-precision observations of the cosmic microwave background, it is also possible to make quantitative comparison of the statistical evidence for or against different inflationary models. This can be done either by comparing the logarithm of the maximum likelihood that can be obtained for the data using each model, i.e., the minimum  $\chi^2$  (with some correction for the number of free parameters in each

model), or by a Bayesian model comparison [83] (see also Sec. 39.3.3 in “Statistics” review).

In such a Bayesian model comparison one computes [7] the evidence,  $\mathcal{E}(\mathcal{D}|\mathcal{M}_A)$  for a model,  $\mathcal{M}_A$ , given the data  $\mathcal{D}$ . This corresponds to the likelihood,  $\mathcal{L}(\theta_{Aj}) = p(\mathcal{D}|\theta_{Aj}, \mathcal{M}_A)$ , integrated over the assumed prior distribution,  $\pi(\theta_{Aj}|\mathcal{M}_A)$ , for all the model parameters  $\theta_{Aj}$ :

$$\mathcal{E}(\mathcal{D}|\mathcal{M}_A) = \int \mathcal{L}(\theta_{Aj}) \pi(\theta_{Aj}|\mathcal{M}_A) d\theta_{Aj}. \quad (23.81)$$

The posterior probability of the model given the data follows from Bayes’ theorem

$$p(\mathcal{M}_A|\mathcal{D}) = \frac{\mathcal{E}(\mathcal{D}|\mathcal{M}_A) \pi(\mathcal{M}_A)}{p(\mathcal{D})}, \quad (23.82)$$

where the prior probability of the model is given by  $\pi(\mathcal{M}_A)$ . Assuming that all models are equally likely a priori,  $\pi(\mathcal{M}_A) = \pi(\mathcal{M}_B)$ , the relative probability of model  $A$  relative to a reference model, in the light of the data, is thus given by the Bayes factor

$$B_{A,\text{ref}} = \frac{\mathcal{E}(\mathcal{D}|\mathcal{M}_A)}{\mathcal{E}(\mathcal{D}|\mathcal{M}_{\text{ref}})}. \quad (23.83)$$

Computation of the multi-dimensional integral (Eq. (23.81)) is a challenging numerical task. Even using an efficient sampling algorithm requires hundreds of thousands of likelihood computations for each model, though slow-roll approximations can be used to calculate rapidly the primordial power spectrum using the APSIC numerical library [7] for a large number of single-field, slow-roll inflation models.

The change in  $\chi^2$  for selected slow-roll models relative to a baseline  $\Lambda$ CDM model is given in Table 1 (taken from [44]). All the inflation models require some amplitude of tensors and so have an increased  $\chi^2$  with respect to the baseline  $\Lambda$ CDM model with a scalar tilt but no tensors. Table Table 23.1 also shows the Bayesian evidence for ( $\ln B_{A,\text{ref}} > 0$ ) or against ( $\ln B_{A,\text{ref}} < 0$ ) a selection of inflation models using the Planck analysis priors [44]. The Starobinsky  $R^2$  inflationary model may be chosen as a reference [44] that provides a good fit to current data. Higgs inflation [58] is indistinguishable using current data, making the model comparison “inconclusive” on the Jeffery’s scale ( $|\ln B_{A,\text{ref}}| < 1$ ). (Recall, though, that this model is disfavoured by the measured values of the Higgs and top quark masses [59].) On the other hand, there is now moderate evidence ( $|\ln B_{A,\text{ref}}| > 2.5$ ) against large-field models such as chaotic inflation with a quadratic potential and strong evidence ( $|\ln B_{A,\text{ref}}| > 5$ ) against chaotic inflation with a quartic potential. Indeed, over 30% of the slow-roll inflation models considered in Ref. [7] are strongly disfavoured by the Planck data.

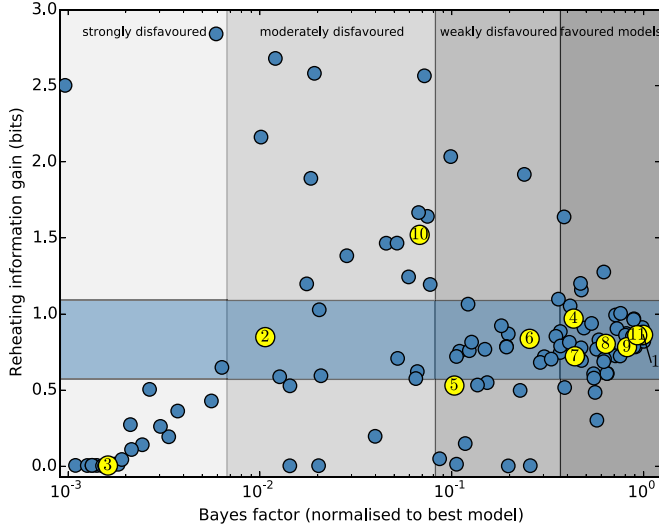
**Table 23.1:** Observational evidence for and against selected inflation models:  $\Delta\chi^2$  is determined relative to a baseline  $\Lambda$ CDM model, and the Bayes factors are calculated relative to Starobinsky  $R^2$  inflation. Results from Planck 2015 analysis [44].

Model	$\Delta\chi^2$	$\ln B_{A,\text{ref}}$
$R^2$ inflation	+0.8	0
Power-law potential $\phi^{2/3}$	+6.5	−2.4
Power-law potential $\phi^2$	+8.6	−4.7
Power-law potential $\phi^4$	+43.3	−23.3
Natural inflation	+7.2	−2.4
SUSY $\alpha$ -attractor	+0.7	−1.8

The Bayes factors for a wide selection of slow-roll inflationary models are displayed in Fig. 23.3, which is adapted from Fig. 3 in [84], where more complete descriptions of the models and the calculations of the Bayes factors are given. Models discussed in this review are highlighted in yellow, and numbered as follows: (1)  $R^2$  inflation



(Sec. 23.4) and models with similar predictions, such as Higgs inflation (Sec. 23.4) and no-scale supergravity inflation (Sec. 23.4); chaotic inflation models (2) with a  $\phi^2$  potential; (3) with a  $\phi^4$  potential; (4) with a  $\phi^{2/3}$  potential, and (5) with a  $\phi^p$  potential marginalising over  $p \in [0.2, 6]$  (Sec. 23.4); hilltop inflation models (6) with  $p = 2$ ; (7) with  $p = 4$  and (8) marginalising over  $p$  (Secinflation:models:hilltop); (9) brane inflation (Secinflation:models:brane); (10) natural inflation (Sec. 23.4); (11) exponential potential models such as  $\alpha$ -attractors (Sec. 23.4). As seen in Fig. 23.3 and discussed in the next Section, constraints on reheating are starting to provide additional information about models of inflation.



**Figure 23.3:** The Bayes factors calculated in [84] for a large sample of inflationary models. Those highlighted in yellow are featured in the this review, according to the numbers listed in the text.

### 23.6. Constraints on Reheating

One connection between inflation and particle physics is provided by inflaton decay, whose products are expected to have thermalized subsequently. As seen in (Eq. (23.55)), the number of e-folds required during inflation depends on details of this reheating process, including the matter density upon reheating, denoted by  $\rho_{\text{th}}$ , which depends in turn on the inflaton decay rate  $\Gamma_\phi$ . We see in Fig. 23.1 that, within any specific inflationary model, both  $n_s$  and particularly  $r$  are sensitive to the value of  $N_*$ . In particular, the one- $\sigma$  uncertainty in the experimental measurement of  $n_s$  is comparable to the variation in many model predictions for  $N_* \in [50, 60]$ . This implies that the data start to constrain scenarios for inflaton decay in many models. For example, it is clear from Fig. 23.1 that  $N_* = 60$  would be preferred over  $N_* = 50$  in a chaotic inflationary model with a quadratic potential.

As a specific example, let us consider  $R^2$  models and related models such as Higgs and no-scale inflation models that predict small values of  $r$  [85]. As seen in Fig. 23.1, within these models the combination of Planck, BICEP2/Keck Array and BAO data would require a limited range of  $n_s$ , corresponding to a limited range of  $N_*$ , as seen by comparing the left and right vertical axes in Fig. 23.4:

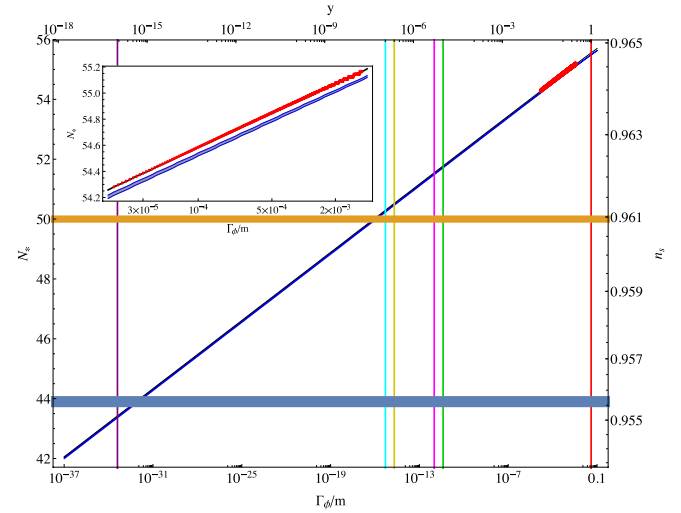
$$N_* \gtrsim 52 \quad (68\% \text{ CL}), \quad N_* \gtrsim 44 \quad (95\% \text{ CL}). \quad (23.84)$$

Within any specific model for inflaton decay, these bounds can be translated into constraints on the effective decay coupling. For example, if one postulates a two-body inflaton decay coupling  $y$ , the bounds (Eq. (23.84)) can be translated into bounds on  $y$ . This is illustrated in Fig. 23.4, where any value of  $N_*$  (on the left vertical

axis), projected onto the diagonal line representing the correlation predicted in  $R^2$ -like models, corresponds to a specific value of the inflaton decay rate  $\Gamma_\phi/m$  (lower horizontal axis) and hence  $y$  (upper horizontal axis):

$$y \gtrsim 10^{-5} \quad (68\% \text{ CL}), \quad y \gtrsim 10^{-15} \quad (95\% \text{ CL}). \quad (23.85)$$

These bounds are not very constraining – although the 68% CL lower bound on  $y$  is already comparable with the electron Yukawa coupling – but can be expected to improve significantly in the coming years and thereby provide significant information on the connections between inflation and particle physics.



**Figure 23.4:** The values of  $N_*$  (left axis) and  $n_s$  (right axis) in  $R^2$  inflation and related models for a wide range of decay rates,  $\Gamma_\phi/m$ , (bottom axis) and corresponding two-body couplings,  $y$  (top axis). The diagonal red line segment shows full numerical results over a restricted range of  $\Gamma_\phi/m$  (which are shown in more detail in the insert), while the diagonal blue strip represents an analytical approximation described in [85]. The difference between these results is indistinguishable in the main plot, but is visible in the insert. The horizontal yellow and blue lines show the 68 and 95% CL lower limits from the Planck 2015 data [44], and the vertical coloured lines correspond to specific models of inflaton decay. Figure taken from [85].

### 23.7. Beyond Single-Field Slow-Roll Inflation

There are numerous possible scenarios beyond the simplest single-field models of slow-roll inflation. These include theories in which non-canonical fields are considered, such as k-inflation [86] or DBI inflation [87], and multiple-field models, such as the curvaton scenario [88]. As well as altering the single-field predictions for the primordial curvature power spectrum (Eq. (23.48)) and the tensor-scalar ratio (Eq. (23.49)), they may introduce new quantities that vanish in single-field slow-roll models, such as isocurvature matter perturbations, corresponding to entropy fluctuations in the photon-to-matter ratio, at first order:

$$S_m = \frac{\delta n_m}{n_m} - \frac{\delta n_\gamma}{n_\gamma} = \frac{\delta \rho_m}{\rho_m} - \frac{3}{4} \frac{\delta \rho_\gamma}{\rho_\gamma}. \quad (23.86)$$

Another possibility is non-Gaussianity in the distribution of the primordial curvature perturbation (see Chap. 28, “Cosmic Microwave Background” review), encoded in higher-order correlators such as the primordial bispectrum [89]

$$\langle \zeta(\mathbf{k}) \zeta(\mathbf{k}') \zeta(\mathbf{k}'') \rangle \equiv (2\pi)^3 \delta(\mathbf{k} + \mathbf{k}' + \mathbf{k}'') B_\zeta(k, k', k''), \quad (23.87)$$

which is often expressed in terms of a dimensionless non-linearity parameter  $f_{\text{NL}} \propto B_\zeta(k, k', k'')/P_\zeta(k)P_\zeta(k')$ . The three-point function (Eq. (23.87)) can be thought of as defined on a triangle whose sides are  $\mathbf{k}, \mathbf{k}', \mathbf{k}''$ , of which only two are independent, since they sum to zero. Further assuming statistical isotropy ensures that the bispectrum depends only on the magnitudes of the three vectors,  $k, k'$  and  $k''$ . The search for  $f_{\text{NL}}$  and other non-Gaussian effects was a prime objective of the Planck data analysis [90].

### 23.7.1. Effective Field Theory of Inflation :

Since slow-roll inflation is a phase of accelerated expansion with an almost constant Hubble parameter, one may think of inflation in terms of an effective theory where the de Sitter spacetime symmetry is spontaneously broken down to RW symmetry by the time-evolution of the Hubble rate,  $\dot{H} \neq 0$ . There is then a Goldstone boson,  $\pi$ , associated with the spontaneous breaking of time-translation invariance, which can be used to study model-independent properties of inflation. The Goldstone boson describes a spacetime-dependent shift of the time coordinate, corresponding to an adiabatic perturbation of the matter fields:

$$\delta\phi_i(t, \vec{x}) = \phi_i(t + \pi(t, \vec{x})) - \phi_i(t). \quad (23.88)$$

Thus adiabatic field fluctuations can be absorbed into the spatial metric perturbation,  $\mathcal{R}$  in Eq. (23.28) at first order, in the comoving gauge:

$$\mathcal{R} = -H\pi, \quad (23.89)$$

where we define  $\pi$  on spatially-flat hypersurfaces. In terms of inflaton field fluctuations, we can identify  $\pi \equiv \delta\phi/\dot{\phi}$ , but in principle this analysis is not restricted to inflation driven by scalar fields.

The low-energy effective action for  $\pi$  can be obtained by writing down the most general Lorentz-invariant action and expanding in terms of  $\pi$ . The second-order effective action for the free-field wave modes,  $\pi_k$ , to leading order in slow roll is then

$$S_\pi^{(2)} = - \int d^4x \sqrt{-g} \frac{M_P^2 \dot{H}}{c_s^2} \left[ \dot{\pi}_k^2 - \frac{c_s^2}{R^2} (\nabla\pi)^2 \right], \quad (23.90)$$

where  $\epsilon_H$  is the Hubble slow-roll parameter (Eq. (23.11)). We identify  $c_s^2$  with an effective sound speed, generalising canonical slow-roll inflation, which is recovered in the limit  $c_s^2 \rightarrow 1$ .

The scalar power spectrum on super-Hubble scales (Eq. (23.48)) is enhanced for a reduced sound speed, leading to a reduced tensor-scalar ratio (Eq. (23.49))

$$\mathcal{P}_\zeta(k) \simeq \frac{4\pi}{M_P^2 c_s^2 \epsilon} \left( \frac{H}{2\pi} \right)^2, \quad r \simeq 16(c_s^2 \epsilon)_*. \quad (23.91)$$

At third perturbative order and to lowest order in derivatives, one obtains [91]

$$S_\pi^{(3)} = \int d^4x \sqrt{-g} \frac{M_P^2 (1 - c_s^2) \dot{H}}{c_s^2} \left[ \frac{\dot{\pi} (\nabla\pi)^2}{R^2} - \left( 1 + \frac{2}{3} \frac{\tilde{c}_3}{c_s^2} \right) \dot{\pi}^3 \right]. \quad (23.92)$$

Note that this expression vanishes for canonical fields with  $c_s^2 = 1$ . For  $c_s^2 \neq 1$  the cubic action is determined by the sound speed and an additional parameter  $\tilde{c}_3$ . Both terms in the cubic action give rise to primordial bispectra that are well approximated by equilateral bispectra. However, the shapes are not identical, so one can find a linear combination for which the equilateral bispectra of each term cancel, giving rise to a distinctive orthogonal-type bispectrum [91].

Analysis based on Planck 2015 temperature and polarisation data has placed bounds on several bispectrum shapes including equilateral and orthogonal shapes [90]:

$$f_{\text{NL}}^{\text{equil}} = -4 \pm 43, \quad f_{\text{NL}}^{\text{orthog}} = -26 \pm 21 \quad (68\% \text{ CL}). \quad (23.93)$$

For the simplest case of a constant sound speed, and marginalising over  $\tilde{c}_3$ , this provides a bound on the inflaton sound speed [90]

$$c_s \geq 0.024 \quad (95\% \text{ CL}). \quad (23.94)$$

For a specific model such as DBI inflation [87], corresponding to  $\tilde{c}_3 = 3(1 - c_s^2)/2$ , one obtains a tighter bound [90]:

$$c_s^{\text{DBI}} \geq 0.087 \quad (95\% \text{ CL}). \quad (23.95)$$

The Planck team have analysed a wide range of non-Gaussian templates from different inflation models, including tests for deviations from an initial Bunch-Davies vacuum state, direction-dependent non-Gaussianity, and feature models with oscillatory bispectra [90]. No individual feature or resonance is above the three- $\sigma$  significance level after accounting for the look-elsewhere effect. These results are consistent with the simplest canonical, slow-roll inflation models, but do not rule out most alternative models; rather, bounds on primordial non-Gaussianity place important constraints on the parameter space for non-canonical models.

### 23.7.2. Multi-Field Fluctuations :

There is a very large literature on two- and multi-field models of inflation, most of which lies beyond the scope of this review [92]. However, two important general topics merit being mentioned here, namely residual isocurvature perturbations and the possibility of non-Gaussian effects in the primordial perturbations.

One might expect that other scalar fields besides the inflaton might have non-negligible values that evolve and fluctuate in parallel with the inflaton, without necessarily making the dominant contribution to the energy density during the inflationary epoch. However, the energy density in such a field might persist beyond the end of inflation before decaying, at which point it might come to dominate (or at least make a non-negligible contribution to) the total energy density. In such a case, its perturbations could end up generating the density perturbations detected in the CMB. This could occur due to a late-decaying scalar field [88] or a field fluctuation that modulates the end of inflation [93] or the inflaton decay [94].

#### 23.7.2.1. Isocurvature Perturbations:

Primordial perturbations arising in single-field slow-roll inflation are necessarily adiabatic, i.e., they affect the overall density without changing the ratios of different contributions, such as the photon-matter ratio,  $\delta(n_\gamma/n_m)/(n_\gamma/n_m)$ . This is because inflaton perturbations represent a local shift of the time, as described in section Sec. 23.7:

$$\pi = \frac{\delta n_\gamma}{\dot{n}_\gamma} = \frac{\delta n_m}{\dot{n}_m}. \quad (23.96)$$

However, any light scalar field (i.e., one with effective mass less than the Hubble scale) acquires a spectrum of nearly scale-invariant perturbations during inflation. Fluctuations orthogonal to the inflaton in field space are decoupled from the inflaton at Hubble-exit, but can affect the subsequent evolution of the density perturbation. In particular, they can give rise to local variations in the equation of state (non-adiabatic pressure perturbations) that can alter the primordial curvature perturbation  $\zeta$  on super-Hubble scales. Since these fluctuations are statistically independent of the inflaton perturbations at leading order in slow-roll [95], non-adiabatic field fluctuations can only increase the scalar power spectrum with respect to adiabatic perturbations at Hubble exit, while leaving the tensor modes unaffected at first perturbative order. Thus the single-field result for the tensor-scalar ratio (Eq. (23.49)) becomes an inequality [96]

$$r \geq 16\epsilon_*. \quad (23.97)$$

Hence an observational upper bound on the tensor-scalar ratio does not bound the slow-roll parameter  $\epsilon$  in multi-field models.

If all the scalar fields present during inflation eventually decay completely into fully thermalized radiation, these field fluctuations are converted fully into adiabatic perturbations in the primordial plasma [97]. On the other hand, non-adiabatic field fluctuations can also leave behind primordial isocurvature perturbations (Eq. (23.86)) after inflation. In multi-field inflation models it is thus possible for non-adiabatic field fluctuations to generate both curvature and isocurvature perturbations leading to correlated primordial perturbations [98].



The amplitudes of any primordial isocurvature perturbations (Eq. (23.86)) are strongly constrained by the current CMB data, especially on large angular scales. Using temperature and low- $\ell$  polarisation data yields the following bound on the amplitude of cold dark matter isocurvature perturbations at scale  $k = 0.002h^{-1}\text{Mpc}^{-1}$  (marginalising over the correlation angle and in the absence of primordial tensor perturbations) [44]:

$$\frac{\mathcal{P}_{S_m}}{\mathcal{P}_\zeta + \mathcal{P}_{S_m}} < 0.020 \text{ at } 95\% \text{ CL}. \quad (23.98)$$

For fully (anti-)correlated isocurvature perturbations, corresponding to a single isocurvature field providing a source for both the curvature and residual isocurvature perturbations, the bounds become significantly tighter [44]:

$$\frac{\mathcal{P}_{S_m}}{\mathcal{P}_\zeta + \mathcal{P}_{S_m}} < 0.0013 \text{ at } 95\% \text{ CL, correlated}, \quad (23.99)$$

$$\frac{\mathcal{P}_{S_m}}{\mathcal{P}_\zeta + \mathcal{P}_{S_m}} < 0.0008 \text{ at } 95\% \text{ CL, anti-correlated}. \quad (23.100)$$

### 23.7.2.2. Local-Type Non-Gaussianity:

Since non-adiabatic field fluctuations in multi-field inflation may lead the evolution of the primordial curvature perturbation at all orders, it becomes possible to generate significant non-Gaussianity in the primordial curvature perturbation. Non-linear evolution on super-Hubble scales leads to local-type non-Gaussianity, where the local integrated expansion is a non-linear function of the local field values during inflation,  $N(\phi_i)$ . While the field fluctuations at Hubble exit,  $\delta\phi_{i*}$ , are Gaussian in the slow-roll limit, the curvature perturbation,  $\zeta = \delta N$ , becomes a non-Gaussian distribution [99]:

$$\zeta = \sum_i \frac{\partial N}{\partial \phi_i} \delta\phi_i + \frac{1}{2} \sum_{i,j} \frac{\partial^2 N}{\partial \phi_i \partial \phi_j} \delta\phi_i \delta\phi_j + \dots \quad (23.101)$$

with non-vanishing bispectrum in the squeezed limit ( $k_1 \approx k_2 \gg k_3$ ):

$$B_\zeta(k_1, k_2, k_3) \approx \frac{12}{5} f_{\text{NL}}^{\text{local}} \frac{\mathcal{P}_\zeta(k_1)}{4\pi k_1^3} \frac{\mathcal{P}_\zeta(k_3)}{4\pi k_3^3}, \quad (23.102)$$

where

$$\frac{6}{5} f_{\text{NL}}^{\text{local}} = \frac{\sum_{i,j} \frac{\partial^2 N}{\partial \phi_i \partial \phi_j}}{\left( \sum_i \frac{\partial N}{\partial \phi_i} \right)^2}. \quad (23.103)$$

Both equilateral and orthogonal bispectra, discussed above in the context of generalised single field inflation, vanish in the squeezed limit, enabling the three types of non-Gaussianity to be distinguished by observations, in principle.

Non-Gaussianity during multi-field inflation is highly model dependent, though  $f_{\text{NL}}^{\text{local}}$  can often be smaller than unity in multi-field slow-roll inflation [100]. Scenarios where a second light field plays a role during or after inflation can make distinctive predictions for  $f_{\text{NL}}^{\text{local}}$ , such as  $f_{\text{NL}}^{\text{local}} = -5/4$  in some curvaton scenarios [99,101] or  $f_{\text{NL}}^{\text{local}} = 5$  in simple modulated reheating scenarios [94,102]. By contrast the constancy of  $\zeta$  on super-Hubble scales in single-field slow-roll inflation leads to a very small non-Gaussianity [103,104], and in the squeezed limit we have the simple result  $f_{\text{NL}}^{\text{local}} = 5(1 - n_s)/12$  [105,106].

A combined analysis of the Planck temperature and polarization data yields the following range for  $f_{\text{NL}}^{\text{local}}$  defined in (Eq. (23.103)):

$$f_{\text{NL}}^{\text{local}} = 0.8 \pm 5.0 \quad (95\% \text{ CL}). \quad (23.104)$$

This sensitivity is sufficient to rule out parameter regimes giving rise to relatively large non-Gaussianity, but insufficient to probe  $f_{\text{NL}}^{\text{local}} = \mathcal{O}(\epsilon)$ , as expected in single-field models, or the range  $f_{\text{NL}}^{\text{local}} = \mathcal{O}(1)$  found in the simplest two-field models.

Local-type primordial non-Gaussianity can also give rise to a striking scale-dependent bias in the distribution of collapsed dark matter halos and thus the galaxy distribution [107,108]. However, bounds from high-redshift galaxy surveys are not yet competitive with the best CMB constraints.

## 23.8. Pre-Inflation and Anomalies in the CMB

Most work on inflation is done in the context of RW cosmology, which already assumes a high degree of symmetry, or small inhomogeneous perturbations (usually first order) about an RW cosmology. The isotropic RW spacetime is an attractor for many homogeneous, but anisotropic cosmologies in the presence of a false vacuum energy density [109] or a scalar field with suitable self-interaction potential energy [110,111]. However it is much harder to establish the range of highly inhomogeneous initial conditions that yield a successful RW Universe, with only limited studies to date (see, e.g., [112,113,114]).

One of the open questions in inflation is the nature of the pre-inflationary state that should have provided suitable initial conditions for inflation. This would need to have satisfied non-trivial homogeneity and isotropy conditions, and one may ask how these could have arisen and whether there may be some observable signature of the pre-inflationary state. In general, one would expect any such effects to appear at large angular scales, i.e., low multipoles  $\ell$ .

Indeed, various anomalies have been noted in the large-scale CMB anisotropies, also discussed in Chap. 28, “Cosmic Microwave Background” review, including a possible suppression of the quadrupole and other very large-scale anisotropies, an apparent feature in the range  $\ell \approx 20$  to 30, and a possible hemispheric asymmetry. None of these are highly statistically significant in view of the limitations due to cosmic variance [44], and they cannot yet be regarded as signatures of some pre-inflationary dynamics such as string theory or the multiverse. However, is a hot topic for present and future analysis.

## 23.9. Prospects for Future Probes of Inflation

When inflation was first proposed [1,2] there was no evidence for the existence of scalar fields or the accelerated expansion of the universe. The situation has changed dramatically in recent years with the observational evidence that the cosmic expansion is currently accelerating and with the discovery of a scalar particle, namely the Higgs boson (see Chap. 11, “Status of Higgs boson physics” review). These discoveries encourage interest in the idea of primordial accelerated expansion driven by a scalar field, i.e., cosmological inflation. In parallel, successive CMB experiments have been consistent with generic predictions of inflationary models, although without yet providing irrefutable evidence.

Prospective future CMB experiments, both ground- and space-based are reviewed in the separate PDG “Cosmic Microwave Background” review, Chap. 28. The main emphasis in CMB experiments in the coming years will be on ground-based experiments providing improved measurements of  $B$ -mode polarization and greater sensitivity to the tensor-to-scalar ratio  $r$ , and more precise measurements at higher  $\ell$  that will constrain  $n_s$  better. As is apparent from Fig. 23.1 and the discussion of models such as  $R^2$  inflation, there is a strong incentive to reach a 5- $\sigma$  sensitivity to  $r \sim 3$  to  $4 \times 10^{-3}$ . This could be achieved with a moderately-sized space mission with large sky coverage [115], improvements in de-lensing and foreground measurements. The discussion in Sec. 23.3 (see also Fig. 23.4), also brought out the importance of reducing the uncertainty in  $n_s$ , as a way to constrain post-inflationary reheating and the connection to particle physics. CMB temperature anisotropies probe primordial density perturbations down to comoving scales of order 50 Mpc, beyond which scale secondary sources of anisotropy dominate. CMB spectral distortions could potentially constrain the amplitude and shape of primordial density perturbations on comoving scales from Mpc to kpc due to distortions caused by the Silk damping of pressure waves in the radiation dominated era, before the last scattering of the CMB photons but after the plasma can be fully thermalised [116].

Improved sensitivity to non-Gaussianities is also a priority. In addition to CMB measurements, future large-scale structure surveys will also have roles to play as probes into models of inflation, for which there are excellent prospects. High-redshift galaxy surveys are sensitive to local-type non-Gaussianity due to the scale-dependent bias induced on large scales. Current surveys such as eBOSS, probing out to redshift  $z \sim 2$ , can reach a precision  $\Delta f_{\text{NL}} \sim 15$ , from measurements of the galaxy power spectrum, or possibly  $\Delta f_{\text{NL}} \sim 10$ ,

if the galaxy bias can be determined independently [117]. Upcoming surveys such as DESI may reach  $\Delta f_{NL} \sim 4$  [118] comparable with the Planck sensitivity. In the future, radio surveys such as SKA will measure large-scale structure out to redshift  $z \sim 3$  [119], initially through mapping the intensity of the neutral hydrogen 21-cm line, and eventually through radio galaxy surveys which will probe local-type non-Gaussianity to  $f_{NL} \sim 1$ .

Galaxy clustering using DESI and Euclid satellite data could also constrain the running of the scalar tilt to a precision of  $\Delta\alpha_s \approx 0.0028$ , a factor of 2 improvement on Planck constraints, or a precision of 0.0016 using LSST data [118].

The proposed SPHEREx satellite mission [120] will use measurements of the galaxy power spectrum to target a measurement of the running of the scalar spectral index with a sensitivity  $\Delta\alpha_s \sim 10^{-3}$  and local-type primordial non-Gaussianity,  $\Delta f_{NL} \sim 1$ . Including information from the galaxy bispectrum one might reduce the measurement error on non-Gaussianity to  $\Delta f_{NL} \sim 0.2$ , making it possible to distinguish between single-field slow-roll models and alternatives such as the curvaton scenario for the origin of structure, which generate  $f_{NL} \sim 1$ .

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## References

1. A.A. Starobinsky, Phys. Lett. **B91**, 99 (1980).
2. A.H. Guth, Phys. Rev. **D23**, 347 (1981).
3. K.A. Olive, Phys. Rept. **190**, 307 (1990).
4. D.H. Lyth and A. Riotto, Phys. Rep. **314**, 1 (1999).
5. A.R. Liddle and D.H. Lyth, *Cosmological inflation and large-scale structure* (Cambridge University Press, 2000).
6. D. Baumann, arXiv:0907.5424 [hep-th].
7. J. Martin, C. Ringeval and V. Vennin, Phys. Dark Univ. **5-6**, 75-235 (2014) [arXiv:1303.3787 [astro-ph.CO]]; J. Martin *et al.*, JCAP **1403** (2014) 039 [arXiv:1312.3529 [astro-ph.CO]]; J. Martin, arXiv:1502.05733 [astro-ph.CO].
8. P.A.R. Ade *et al.* [Planck Collab.], arXiv:1502.01589 [astro-ph.CO].
9. A.H. Guth and E.J. Weinberg, Nucl. Phys. B **212**, 321 (1983).
10. D. La and P.J. Steinhardt, Phys. Rev. Lett. **62**, 376 (1989) [Phys. Rev. Lett. **62**, 1066 (1989)].
11. A.D. Linde, Phys. Lett. **B249**, 18 (1990).
12. F.C. Adams and K. Freese, Phys. Rev. **D43**, 353 (1991) [hep-ph/0504135].
13. A.D. Linde, Phys. Lett. **B108**, 389 (1982).
14. A. Albrecht and P.J. Steinhardt, Phys. Rev. Lett. **48**, 1220 (1982).
15. W.H. Press, Phys. Scr. **21** (1980) 702; S.W. Hawking, Phys. Lett. **115B** (1982) 295; A.A. Starobinsky, Phys. Lett. **117B** (1982) 175; A.H. Guth and S.Y. Pi, Phys. Rev. Lett. **49** (1982) 1110.
16. J.M. Bardeen, P.J. Steinhardt and M.S. Turner, Phys. Rev. **D28**, 679 (1983).
17. V.F. Mukhanov and G.V. Chibisov, JETP Lett. **33** (1981) 532.
18. K.S. Stelle, Gen. Rel. Grav. **9** (1978) 353.
19. B. Whitt, Phys. Lett. **B145**, 176 (1984).
20. D. Wands, Class. Quant. Grav. **11**, 269 (1994) [gr-qc/9307034].
21. S.R. Coleman and F. De Luccia, Phys. Rev. **D21**, 3305 (1980).
22. M. Sasaki *et al.*, Phys. Lett. **B317**, 510 (1993).
23. M. Bucher, A.S. Goldhaber and N. Turok, Phys. Rev. **D52**, 3314 (1995) [hep-ph/9411206].
24. A.D. Linde and A. Mezhlumian, Phys. Rev. **D52**, 6789 (1995) [astro-ph/9506017].
25. A.R. Liddle, P. Parsons and J.D. Barrow, Phys. Rev. **D50**, 7222 (1994) [astro-ph/9408015].
26. L. Kofman, A.D. Linde and A.A. Starobinsky, Phys. Rev. **D56**, 3258 (1997) [hep-ph/9704452].
27. B.A. Bassett, S. Tsujikawa and D. Wands, Rev. Mod. Phys. **78**, 537 (2006) [astro-ph/0507632].
28. A.D. Linde, Phys. Rev. **D49**, 748 (1994) [astro-ph/9307002].
29. A.D. Dolgov and A.D. Linde, Phys. Lett. **B116**, 329 (1982).
30. J.R. Ellis *et al.*, Nucl. Phys. B **238**, 453 (1984).
31. M. Kawasaki and T. Moroi, Prog. Theor. Phys. **93**, 879 (1995) [hep-ph/9403364, hep-ph/9403061].
32. J.H. Traschen and R.H. Brandenberger, Phys. Rev. **D42**, 2491 (1990).
33. G.V. Chibisov and V.F. Mukhanov, Mon. Not. Roy. Astron. Soc. **200**, 535 (1982).
34. H. Kodama and M. Sasaki, Prog. Theor. Phys. Suppl. **78**, 1 (1984).
35. V.F. Mukhanov, H.A. Feldman and R.H. Brandenberger, Phys. Rept. **215**, 203 (1992).
36. K.A. Malik and D. Wands, Phys. Rept. **475**, 1 (2009) [arXiv:0809.4944 [astro-ph]].
37. M. Sasaki, Prog. Theor. Phys. **76**, 1036 (1986) V.F. Mukhanov, Sov. Phys. JETP **67**, 1297 (1988) [Zh. Eksp. Teor. Fiz. **94N7**, 1 (1988)].
38. A.A. Starobinsky, JETP Lett. **30**, 682 (1979) [Pisma Zh. Eksp. Teor. Fiz. **30**, 719 (1979)].
39. E.D. Stewart and D.H. Lyth, Phys. Lett. **B302**, 171 (1993) [gr-qc/9302019].
40. D. Wands *et al.*, Phys. Rev. **D62**, 043527 (2000) [astro-ph/0003278].
41. A.R. Liddle and S.M. Leach, Phys. Rev. **D68**, 103503 (2003) [astro-ph/0305263].
42. P.A.R. Ade *et al.* [BICEP2/Keck and Planck Collabs.], Phys. Rev. Lett. **114**, 101301 (2015) [arXiv:1502.00612 [astro-ph.CO]]; P.A.R. Ade *et al.* [BICEP2 and Keck Array Collabs.], Phys. Rev. Lett. **116**, 031302 (2016) [arXiv:1510.09217 [astro-ph.CO]].
43. S.M. Leach, A.R. Liddle, J. Martin, and D.J. Schwarz, Phys. Rev. **D66**, 23515 (2002).
44. P.A.R. Ade *et al.* [Planck Collab.], arXiv:1502.02114 [astro-ph.CO].
45. D. Kazanas, Astrophys. J. **241** (1980) L59.
46. K. Sato, Mon. Not. Roy. Astron. Soc. **195**, 467 (1981).
47. A. Billoire and K. Tamvakis, Nucl. Phys. B **200** (1982) 329; J.D. Breit, S. Gupta and A. Zaks, Phys. Rev. Lett. **51**, 1007 (1983).
48. A.D. Linde, Phys. Lett. **B129**, 177 (1983).
49. D.V. Nanopoulos, K.A. Olive and M. Srednicki, Phys. Lett. **B127**, 30 (1983).
50. J. Ellis, D.V. Nanopoulos and K.A. Olive, Phys. Rev. Lett. **111**, 111301 (2013) [Phys. Rev. Lett. **111**, 129902 (2013)] [arXiv:1305.1247 [hep-th]].
51. C. Pallis and Q. Shafi, JCAP **1503** (2015) 03, 023 [arXiv:1412.3757 [hep-ph]].
52. L. Boubekeur and D.H. Lyth, JCAP **0507** (2005) 010 [hep-ph/0502047].
53. G.R. Dvali, Q. Shafi and S. Solganik, hep-th/0105203; J. Garcia-Bellido, R. Rabadan and F. Zamora, JHEP **0201** (2002) 036 [hep-th/0112147]; S. Kachru *et al.*, JCAP **0310** (2003) 013 [hep-th/0308055].
54. K. Freese, J.A. Frieman and A.V. Olinto, Phys. Rev. Lett. **65**, 3233 (1990) F.C. Adams *et al.*, Phys. Rev. **D47**, 426 (1993) [hep-ph/9207245].
55. For a recent review of axion inflation, see E. Pajer and M. Peloso, Class. Quant. Grav. **30** (2013) 214002 [arXiv:1305.3557 [hep-th]].
56. E. Silverstein and A. Westphal, Phys. Rev. **D78**, 106003 (2008) [arXiv:0803.3085 [hep-th]].
57. L. McAllister, E. Silverstein and A. Westphal, Phys. Rev. **D82**, 046003 (2010) [arXiv:0808.0706 [hep-th]].
58. F.L. Bezrukov and M. Shaposhnikov, Phys. Lett. **B659**, 703 (2008) [arXiv:0710.3755 [hep-th]].
59. D. Buttazzo *et al.*, JHEP **1312** (2013) 089 [arXiv:1307.3536 [hep-ph]].

60. H.P. Nilles, Phys. Rept. **110** (1984) 1; H.E. Haber and G.L. Kane, Phys. Rept. **117** (1985) 75.
61. J.R. Ellis *et al.*, Phys. Lett. **B118**, 336 (1982).
62. N. Okada and Q. Shafi, arXiv:1311.0921 [hep-ph].
63. H. Murayama *et al.*, Phys. Rev. Lett. **70**, 1912 (1993).
64. J. Ellis, M. Fairbairn and M. Sueiro, JCAP **1402** (2014) 044 [arXiv:1312.1353 [astro-ph.CO]].
65. D. Croon, J. Ellis and N.E. Mavromatos, Phys. Lett. **B724**, 165 (2013) [arXiv:1303.6253 [astro-ph.CO]].
66. G.R. Dvali, Q. Shafi and R.K. Schaefer, Phys. Rev. Lett. **73**, 1886 (1994) [hep-ph/9406319].
67. D.V. Nanopoulos *et al.*, Phys. Lett. **B123**, 41 (1983).
68. A.B. Goncharov and A.D. Linde, Phys. Lett. **B139**, 27 (1984).
69. E. Cremmer *et al.*, Nucl. Phys. B **212** (1983) 413.
70. R. Holman, P. Ramond and G.G. Ross, Phys. Lett. **B137**, 343 (1984).
71. E.J. Copeland *et al.*, Phys. Rev. **D49**, 6410 (1994) [astro-ph/9401011]; E.D. Stewart, Phys. Rev. **D51**, 6847 (1995) [hep-ph/9405389].
72. G.D. Coughlan *et al.*, Phys. Lett. B **131**, 59 (1983); A.S. Goncharov, A.D. Linde and M.I. Vysotsky, Phys. Lett. **B147**, 279 (1984); T. Banks, D.B. Kaplan and A.E. Nelson, Phys. Rev. **D49**, 779 (1994) [hep-ph/9308292]; B. De Carlos *et al.*, Phys. Lett. **B318**, 447 (1993) [hep-ph/9308325]; M. Kawasaki, T. Moroi and T. Yanagida Phys. Lett. **B370**, 52 (1996) [hep-ph/9509399].
73. J. Ellis, D.V. Nanopoulos, and M. Quiros, Phys. Lett. **B174**, 176 (1986).
74. T. Moroi, M. Yamaguchi and T. Yanagida Phys. Lett. **B342**, 105 (1995) [hep-ph/9409367].
75. M. Kawasaki, M. Yamaguchi and T. Yanagida, Phys. Rev. Lett. **85**, 3572 (2000) [hep-ph/0004243]; K. Nakayama, F. Takahashi and T.T. Yanagida, JCAP **1308** (2013) 038 [arXiv:1305.5099 [hep-ph]].
76. E. Cremmer *et al.*, Phys. Lett. **B133**, 61 (1983).
77. For some older examples of no-scale inflationary models see, e.g., A.S. Goncharov and A.D. Linde, Class. Quant. Grav. **1**, L75 (1984); C. Kounnas and M. Quiros, Phys. Lett. **B151**, 189 (1985); J.R. Ellis *et al.*, Phys. Lett. **B152**, 175 (1985) [Erratum-ibid. **156B** (1985) 452].
78. J.R. Ellis, C. Kounnas and D.V. Nanopoulos, Nucl. Phys. B **247** (1984) 373.
79. E. Witten, Phys. Lett. **B155**, 151 (1985).
80. J. Ellis, D.V. Nanopoulos and K.A. Olive, JCAP **1310** (2013) 009 [arXiv:1307.3537].
81. For a recent review and more references, see J. Ellis *et al.*, arXiv:1507.02308 [hep-ph].
82. R. Kallosh, A. Linde and D. Roest, JHEP **1311** (2013) 198 [arXiv:1311.0472].
83. A.R. Liddle, Mon. Not. Roy. Astron. Soc. **377**, L74 (2007) [astro-ph/0701113].
84. J. Martin, C. Ringeval and V. Vennin, arXiv:1603.02606 [astro-ph.CO] We thank Vincent Vennin for providing the adapted version of Fig. 3 from this reference.
85. J. Ellis *et al.*, JCAP **1507** (2015) 07, 050 [arXiv:1505.06986 [hep-ph]].
86. C. Armendariz-Picon, T. Damour and V.F. Mukhanov, Phys. Lett. **B458**, 209 (1999) [hep-th/9904075].
87. M. Alishahiha, E. Silverstein and D. Tong, Phys. Rev. **D70**, 123505 (2004) [hep-th/0404084].
88. K. Enqvist and M.S. Sloth, Nucl. Phys. B **626** (2002) 395 [hep-ph/0109214]; D.H. Lyth and D. Wands, Phys. Lett. **B524**, 5 (2002) [hep-ph/0110002]; T. Moroi and T. Takahashi, Phys. Lett. **B522**, 215 (2001) [Phys. Lett. **B539**, 303 (2002)] [hep-ph/0110096].
89. N. Bartolo *et al.*, Phys. Rept. **402**, 103 (2004) [astro-ph/0406398].
90. P.A.R. Ade *et al.* [Planck Collab.], arXiv:1502.01592 [astro-ph.CO].
91. L. Senatore, K.M. Smith and M. Zaldarriaga, JCAP **1001**, 028 (2010) [arXiv:0905.3746 [astro-ph.CO]].
92. See, for example, C. Gordon *et al.*, Phys. Rev. **D63**, 023506 (2001) [astro-ph/0009131]; C.T. Byrnes and D. Wands, Phys. Rev. **D74**, 043529 (2006) [astro-ph/0605679]; R. Easther *et al.*, Phys. Rev. Lett. **112**, 161302 (2014) [arXiv:1312.4035 [astro-ph.CO]]; J. Ellis *et al.*, JCAP **1501** (2015) 01, 010 [arXiv:1409.8197 [hep-ph]]; S. Renaux-Petel and K. Turzynski, JCAP **1506** (2015) 06, 010 [arXiv:1405.6195 [astro-ph.CO]].
93. D.H. Lyth, JCAP **0511**, 006 (2005) [astro-ph/0510443].
94. G. Dvali, A. Gruzinov and M. Zaldarriaga, Phys. Rev. **D69**, 023505 (2004) [astro-ph/0303591].
95. C.T. Byrnes and D. Wands, Phys. Rev. **D74**, 043529 (2006) [astro-ph/0605679].
96. D. Wands *et al.*, Phys. Rev. **D66**, 043520 (2002) [astro-ph/0205253].
97. S. Weinberg, Phys. Rev. **D67**, 123504 (2003) [astro-ph/0302326].
98. D. Langlois, Phys. Rev. **D59**, 123512 (1999) [astro-ph/9906080].
99. D.H. Lyth and Y. Rodriguez, Phys. Rev. Lett. **95**, 121302 (2005) [astro-ph/0504045].
100. F. Vernizzi and D. Wands, JCAP **0605**, 019 (2006) [astro-ph/0603799].
101. M. Sasaki, J. Valiviita and D. Wands, Phys. Rev. **D74**, 103003 (2006) [astro-ph/0607627].
102. G. Dvali, A. Gruzinov and M. Zaldarriaga, Phys. Rev. **D69**, 083505 (2004) [astro-ph/0305548].
103. D.S. Salopek and J.R. Bond, Phys. Rev. **D43**, 1005 (1991).
104. A. Gangui *et al.*, Astrophys. J. **430**, 447 (1994) [astro-ph/9312033].
105. J.M. Maldacena, JHEP **0305**, 013 (2003) [astro-ph/0210603].
106. V. Acquaviva *et al.*, Nucl. Phys. B **667**, 119 (2003) [astro-ph/0209156].
107. N. Dalal *et al.*, Phys. Rev. **D77**, 123514 (2008) [arXiv:0710.4560 [astro-ph]].
108. S. Matarrese and L. Verde, Astrophys. J. **677**, L77 (2008) [arXiv:0801.4826 [astro-ph]].
109. R.M. Wald, Phys. Rev. **D28**, 2118 (1983).
110. M. Heusler, Phys. Lett. **B253**, 33 (1991).
111. Y. Kitada and K.-i. Maeda, Phys. Rev. **D45**, 1416 (1992).
112. D.S. Goldwirth and T. Piran, Phys. Rept. **214**, 223 (1992).
113. T. Vachaspati and M. Trodden, Phys. Rev. **D61**, 023502 (1999) [gr-qc/9811037].
114. W.E. East *et al.*, arXiv:1511.05143 [hep-th].
115. Proposed space missions include PIXIE: A. Kogut *et al.*, JCAP **1107** (2011) 025 [arXiv:1105.2044 [astro-ph.CO]]; LiteBird: <http://litebird.jp/eng/> and Cosmic Origins Explorer (CORe): <http://www.core-mission.org>.
116. J. Chluba, J. Hamann and S.P. Patil, Int. J. Mod. Phys. D **24**, no. 10, 1530023 (2015) [arXiv:1505.01834 [astro-ph.CO]].
117. G.B. Zhao *et al.*, Mon. Not. Roy. Astron. Soc. **457**, 2377 (2016) [arXiv:1510.08216 [astro-ph.CO]].
118. A. Font-Ribera *et al.*, JCAP **1405**, 023 (2014) [arXiv:1308.4164 [astro-ph.CO]].
119. R. Maartens *et al.* [SKA Cosmology SWG Collab.], PoS AASKA **14**, 016 (2015) [arXiv:1501.04076 [astro-ph.CO]].
120. O. Dor *et al.*, arXiv:1412.4872 [astro-ph.CO].

## 24. BIG-BANG NUCLEOSYNTHESIS

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Big-Bang nucleosynthesis (BBN) offers the deepest reliable probe of the early Universe, being based on well-understood Standard Model physics [1]. Predictions of the abundances of the light elements, D,  $^3\text{He}$ ,  $^4\text{He}$ , and  $^7\text{Li}$ , synthesized at the end of the ‘first three minutes’, are in good overall agreement with the primordial abundances inferred from observational data, thus validating the standard hot Big-Bang cosmology (see [2–5] for reviews). This is particularly impressive given that these abundances span nine orders of magnitude – from  $^4\text{He}/\text{H} \sim 0.08$  down to  $^7\text{Li}/\text{H} \sim 10^{-10}$  (ratios by number). Thus BBN provides powerful constraints on possible deviations from the standard cosmology, and on new physics beyond the Standard Model [6–9].

## 24.1. Theory

The synthesis of the light elements is sensitive to physical conditions in the early radiation-dominated era at a temperature  $T \sim 1$  MeV, corresponding to an age  $t \sim 1$  s. At higher temperatures, weak interactions were in thermal equilibrium, thus fixing the ratio of the neutron and proton number densities to be  $n/p = e^{-Q/T}$ , where  $Q = 1.293$  MeV is the neutron-proton mass difference. As the temperature dropped, the neutron-proton inter-conversion rate per nucleon,  $\Gamma_{n \leftrightarrow p} \sim G_F^2 T^5$ , fell faster than the Hubble expansion rate,  $H \sim \sqrt{g_* G_N} T^2$ , where  $g_*$  counts the number of relativistic particle species determining the energy density in radiation (see ‘Big Bang Cosmology’ review). This resulted in departure from chemical equilibrium (‘freeze-out’) at  $T_{\text{fr}} \sim (g_* G_N / G_F^4)^{1/6} \simeq 1$  MeV. The neutron fraction at this time,  $n/p = e^{-Q/T_{\text{fr}}} \simeq 1/6$ , is thus sensitive to every known physical interaction, since  $Q$  is determined by both strong and electromagnetic interactions while  $T_{\text{fr}}$  depends on the weak as well as gravitational interactions. Moreover, the sensitivity to the Hubble expansion rate affords a probe of, *e.g.*, the number of relativistic neutrino species [10]. After freeze-out, the neutrons were free to  $\beta$ -decay, so the neutron fraction dropped to  $n/p \simeq 1/7$  by the time nuclear reactions began. A simplified analytic model of freeze-out yields the  $n/p$  ratio to an accuracy of  $\sim 1\%$  [11,12].

The rates of these reactions depend on the density of baryons (strictly speaking, nucleons), which is usually expressed normalized to the relic blackbody photon density as  $\eta \equiv n_b/n_\gamma$ . As we shall see, all the light-element abundances can be explained with  $\eta_{10} \equiv \eta \times 10^{10}$  in the range 5.8–6.6 (95% CL). With  $n_\gamma$  fixed by the present CMB temperature 2.7255 K (see ‘Cosmic Microwave Background’ review), this can be stated as the allowed range for the baryon mass density today,  $\rho_b = (3.9\text{--}4.6) \times 10^{-31} \text{ g cm}^{-3}$ , or as the baryonic fraction of the critical density,  $\Omega_b = \rho_b/\rho_{\text{crit}} \simeq \eta_{10} h^{-2}/274 = (0.021\text{--}0.024)h^{-2}$ , where  $h \equiv H_0/100 \text{ km s}^{-1} \text{ Mpc}^{-1}$  is the present Hubble parameter (see Cosmological Parameters review).

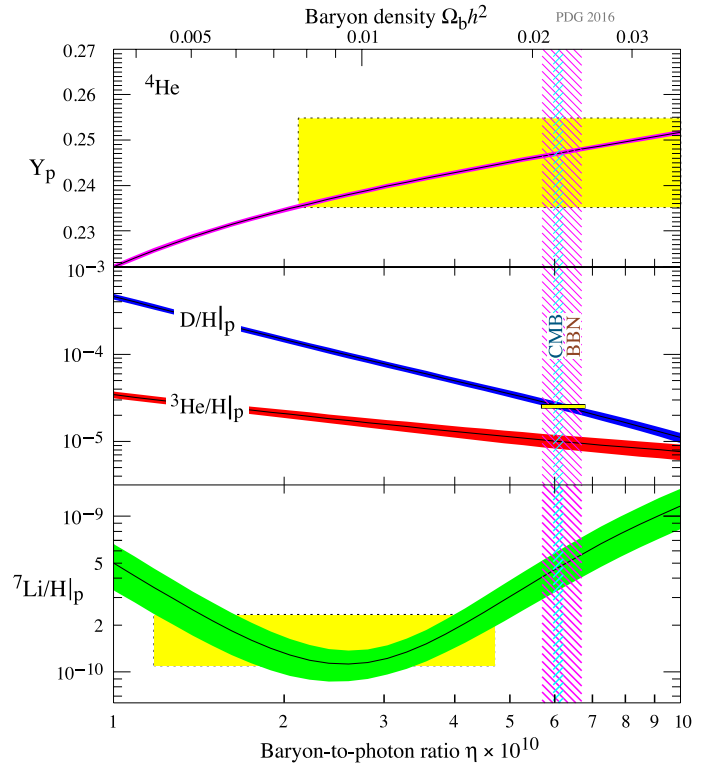
The nucleosynthesis chain begins with the formation of deuterium in the process  $p(n, \gamma)\text{D}$ . However, photo-dissociation by the high number density of photons delays production of deuterium (and other complex nuclei) until well after  $T$  drops below the binding energy of deuterium,  $\Delta_D = 2.23$  MeV. The quantity  $\eta^{-1}e^{-\Delta_D/T}$ , *i.e.*, the number of photons per baryon above the deuterium photo-dissociation threshold, falls below unity at  $T \simeq 0.1$  MeV; nuclei can then begin to form without being immediately photo-dissociated again. Only 2-body reactions, such as  $\text{D}(p, \gamma)^3\text{He}$ ,  $^3\text{He}(\text{D}, p)^4\text{He}$ , are important because the density by this time has become rather low – comparable to that of air!

Nearly all neutrons end up bound in the most stable light element  $^4\text{He}$ . Heavier nuclei do not form in any significant quantity both because of the absence of stable nuclei with mass number 5 or 8 (which impedes nucleosynthesis via  $n^4\text{He}$ ,  $p^4\text{He}$  or  $^4\text{He}^4\text{He}$  reactions), and the large Coulomb barriers for reactions such as  $^3\text{He}(^4\text{He}, \gamma)^7\text{Li}$  and  $^3\text{He}(^4\text{He}, \gamma)^7\text{Be}$ . Hence the primordial mass fraction of  $^4\text{He}$ ,  $Y_p \equiv \rho(^4\text{He})/\rho_b$ , can be estimated by the simple counting argument

$$Y_p = \frac{2(n/p)}{1 + n/p} \simeq 0.25. \quad (24.1)$$

There is little sensitivity here to the actual nuclear reaction rates, which are, however, important in determining the other ‘left-over’ abundances: D and  $^3\text{He}$  at the level of a few times  $10^{-5}$  by number relative to H, and  $^7\text{Li}/\text{H}$  at the level of about  $10^{-10}$  (when  $\eta_{10}$  is in the range 1–10). These values can be understood in terms of approximate analytic arguments [12,13]. The experimental parameter most important in determining  $Y_p$  is the neutron lifetime,  $\tau_n$ , which normalizes (the inverse of)  $\Gamma_{n \leftrightarrow p}$ . Its value has recently been significantly revised downwards to  $\tau_n = 880.3 \pm 1.1$  s (see *N Baryons* Listing).

The elemental abundances shown in Fig. 24.1 as a function of  $\eta_{10}$  were calculated [14] using an updated version [15] of the Wagoner code [1]; other versions [16–18] are publicly available. The  $^4\text{He}$  curve includes small corrections due to radiative processes at zero and finite temperatures [19], non-equilibrium neutrino heating during  $e^\pm$  annihilation [20], and finite nucleon mass effects [21]; the range reflects primarily the  $2\sigma$  uncertainty in the neutron lifetime. The spread in the curves for D,  $^3\text{He}$ , and  $^7\text{Li}$  corresponds to the  $2\sigma$  uncertainties in nuclear cross sections, as estimated by Monte Carlo methods [15,22–24]. The input nuclear data have been carefully reassessed [14, 24–28], leading to improved precision in the abundance predictions. In particular, the uncertainty in  $^7\text{Li}/\text{H}$  at interesting values of  $\eta$  has been reduced recently by a factor  $\sim 2$ , a consequence of a similar reduction in the error budget [29] for the dominant mass-7 production channel  $^3\text{He}(^4\text{He}, \gamma)^7\text{Be}$ . Polynomial fits to the predicted abundances and the error correlation matrix have been given [23,30]. The boxes in Fig. 24.1 show the observationally inferred primordial abundances with their associated uncertainties, as discussed below.



**Figure 24.1:** The primordial abundances of  $^4\text{He}$ , D,  $^3\text{He}$ , and  $^7\text{Li}$  as predicted by the standard model of Big-Bang nucleosynthesis—the bands show the 95% CL range [5]. Boxes indicate the observed light element abundances. The narrow vertical band indicates the CMB measure of the cosmic baryon density, while the wider band indicates the BBN concordance range (both at 95% CL).

The nuclear reaction cross sections important for BBN have all been measured at the relevant energies. We will see, however, that recently there have been substantial advances in the precision of light

element observations (e.g., D/H) and in cosmological parameters (e.g., from *Planck*). This motivates corresponding improvement in BBN precision and thus in the key reaction cross sections. For example, it has been suggested [31] that  $d(p,\gamma)^3\text{He}$  measurements may suffer from systematic errors and be inferior to *ab initio* theory; if so, this could alter D/H abundances at a level that is now significant.

## 24.2. Light Element Abundances

BBN theory predicts the universal abundances of D,  $^3\text{He}$ ,  $^4\text{He}$ , and  $^7\text{Li}$ , which are essentially fixed by  $t \sim 180$  s. Abundances are, however, observed at much later epochs, after stellar nucleosynthesis has commenced. This produces heavy elements such as C, N, O, and Fe (‘metals’), while the ejected remains of this stellar processing alters the light element abundances from their primordial values. Thus, one seeks astrophysical sites with low metal abundances, in order to measure light element abundances that are closer to primordial. For all of the light elements, systematic errors are the dominant limitation to the precision with which primordial abundances can be inferred.

BBN is the only significant source of deuterium, which is entirely destroyed when it is cycled into stars [32]. Thus, any detection provides a lower limit to primordial D/H, and an upper limit on  $\eta_{10}$ ; for example, the local interstellar value of  $\text{D/H} = (1.56 \pm 0.40) \times 10^{-5}$  [33] requires  $\eta_{10} \leq 9$ . The best proxy to the primordial value of D is its measure in distant and chemically unprocessed matter where stellar processing (astration) is minimal [32]. This has become possible with the advent of large telescopes, but after two decades of observational efforts we have only about a dozen determinations [34–43]. High-resolution spectra reveal the presence of D in high-redshift, low-metallicity quasar absorption systems via its isotope-shifted Lyman- $\alpha$  absorption features; these are, unfortunately, usually obscured by the Lyman- $\alpha$  forest. The available D measurements are performed in systems with metallicities from 0.1 to 0.001 Solar where no significant astration is expected [35]. In the best-measured systems, D/H shows no hint of correlation with metallicity, redshift or the hydrogen column density  $N(\text{H}) (= \int_{\text{los}} n_{\text{H}} ds)$  integrated over the line-of-sight through the absorber. This is consistent with the measured D/H being representative of the primordial value.

The first measurements in ‘damped’ Lyman- $\alpha$  systems (DLAs:  $N(\text{H}) > 10^{20} \text{ cm}^{-2}$ ) [34,36] showed that D/H can be measured in this class of absorbers where the Lorentzian damping wings of Lyman- $\alpha$  and Lyman- $\beta$  (if relatively uncontaminated by Lyman- $\alpha$  clouds) provide a precise H column density. Subsequently DLA systems have been found that also show resolved higher members of the Lyman series. Systems with a particularly simple kinematic structure are desirable to avoid uncertainties with complex, only partially resolved components. Recently a DLA showing 13 resolved D I absorption lines has been analyzed together with 4 other suitable systems. This provides a strikingly improved precision over earlier work, with a weighted mean of  $\log(\text{D/H}) = -4.597 \pm 0.006$ , corresponding to [37]

$$\text{D/H}|_{\text{p}} = (2.53 \pm 0.04) \times 10^{-5}. \quad (24.2)$$

D/H values in the Galaxy show an unexpected scatter of a factor of  $\sim 2$  [44], with a bimodal distribution as well as an anti-correlation with metal abundances. This suggests that interstellar D not only suffers stellar astration but also partly resides in dust particles that evade gas-phase observations. This is supported by a measurement in the lower halo [45], which indicates that the Galactic D abundance has decreased by a factor of only  $1.1 \pm 0.13$  since its formation. However in the DLA the dust content is apparently quite small; this is implied by the abundances of refractory elements such as Fe, Cr and Si, which are in nearly Solar proportions. Thus, the value derived in Eq. (24.2) appears safe against D depletion into dust grains.

The primordial  $^4\text{He}$  abundance is best determined through recombination emission lines of He and H in the most metal-poor extragalactic H II (ionized) regions, *viz.* blue compact galaxies. There is now a large body of data on  $^4\text{He}$  and CNO in these galaxies, with over 1000 such systems in the Sloan Digital Sky Survey alone [46,53]. These data confirm that the small stellar contribution to the

helium abundance is positively correlated with metal production, so extrapolation to zero metallicity gives the primordial  $^4\text{He}$  abundance  $Y_{\text{p}}$ . However, H II regions are complex systems and several physical parameters enter in the He/H determination, notably the electron density and temperature, as well as reddening. Thus systematic effects dominate the uncertainties in the abundance determination [46,47]. A major step forward has been the inclusion of the He  $\lambda 10830$  infrared emission line which shows a strong dependence on the electron density and is thus useful to break the degeneracy with the temperature, allowing for a more robust helium abundance determination. In recent work that has accounted for the underlying  $^4\text{He}$  stellar absorption, and/or the newly derived values of the HeI-recombination and H-excitation-collisional coefficients, the  $^4\text{He}$  abundances have increased significantly. Two recent results are  $Y_{\text{p}} = 0.2449 \pm 0.0040$  [51] and  $Y_{\text{p}} = 0.2551 \pm 0.0022$  [52] (see Ref. [53] and references therein for previous determinations). Our recommended  $^4\text{He}$  abundance is

$$Y_{\text{p}} = 0.245 \pm 0.004, \quad (24.3)$$

but the matter is far from settled given the two measurements are only marginally consistent.

As we will see in more detail below, the primordial abundance of  $^7\text{Li}$  now plays a central role in BBN, and possibly points to new physics. The systems best suited for Li observations are metal-poor (Pop II) stars in the spheroid of the Galaxy, which have metallicities going down to perhaps  $10^{-5}$  of the solar value [56]. Observations have long shown [57–60] that Li does not vary significantly in Pop II stars with metallicities  $\lesssim 1/30$  of Solar — the ‘Spite plateau’ [57]. However there are systematic uncertainties due to different techniques used to determine the physical parameters (*e.g.*, the temperature) of the stellar atmosphere in which the Li absorption line is formed. Different analyses and in some cases different stars and stellar systems (globular clusters), yield  $\text{Li/H}|_{\text{p}} = (1.7 \pm 0.3) \times 10^{-10}$  [60],  $\text{Li/H}|_{\text{p}} = (2.19 \pm 0.28) \times 10^{-10}$  [61], and  $\text{Li/H}|_{\text{p}} = (1.86 \pm 0.23) \times 10^{-10}$  [62].

Recent observations find a puzzling drop in Li/H in metal-poor stars with  $[\text{Fe/H}] \equiv \log_{10}[(\text{Fe/H})/(\text{Fe/H})_{\odot}] < -3.0$  [63,64,65] particularly at the very low metallicity end. Li is not detected at all, or is well below than the Spite Plateau, in *all* the 5 extremely metal poor dwarfs with metallicities  $[\text{Fe/H}] \lesssim -4.5$ , where it ought to be present. The reason is not known and the same effect(s) may also produce the ‘melting’ of the Li plateau at metallicities  $[\text{Fe/H}] \approx -3.0$  [64,65] thus making quite uncertain any primordial Li value extracted by extrapolating to zero metallicity. To estimate the primordial value it is therefore safer to consider stars with  $-2.8 < [\text{Fe/H}] < -1.5$  [65], which yields

$$\text{Li/H}|_{\text{p}} = (1.6 \pm 0.3) \times 10^{-10}. \quad (24.4)$$

However, the evidence that something is depleting Li at the low metallicity end suggests that its abundance may also be modified in halo stars with moderate metallicity so the observed abundance should be considered a lower bound rather than a measure of the primordial Li. In fact Li in Pop II stars may have been partially destroyed due to mixing of the outer layers with the hotter interior [68]. Such processes can be constrained by the absence of significant scatter in Li versus Fe [59] but Li depletion by a factor as large as  $\sim 1.8$  has been suggested [69]. A new model [70] predicts Li significantly destroyed in the pre-MS phase by overshoot mixing, and then partially restored by late accretion of fresh non-Li depleted material has been proposed. They show that it is also possible to recover the Spite plateau while starting from an initial  $\text{Li/H}|_{\text{p}} = 5.3 \times 10^{-10}$  suggested by both Planck and Deuterium baryonic density estimation [67,70].

Stellar determination of Li abundances typically sum over both stable isotopes  $^6\text{Li}$  and  $^7\text{Li}$ . Recent high-precision measurements are sensitive to the tiny isotopic shift in Li absorption (which manifests itself in the shape of the blended, thermally broadened line) and indicate  $^6\text{Li}/^7\text{Li} \leq 0.05$  [71,72], thus confirming that  $^7\text{Li}$  is dominant. A claim of a  $^6\text{Li}$  plateau (analogous to the  $^7\text{Li}$  plateau) has been made [71], suggesting a significant primordial  $^6\text{Li}$  abundance. This has, however, been challenged by new observations and analyses [73,74,72], which show that stellar convective motions can generate asymmetries

in the line shape that mimic the presence of  ${}^6\text{Li}$ . Hence the deduced abundance ratio  ${}^6\text{Li}/{}^7\text{Li} < 0.05$  in the best studied stars presently provides a robust upper limit on the  ${}^6\text{Li}$  abundance [72].

Turning to  ${}^3\text{He}$ , the only data available are from the Solar system and (high-metallicity) H II regions in our Galaxy [75]. This makes inferring the primordial abundance difficult, a problem compounded by the fact that stellar nucleosynthesis models for  ${}^3\text{He}$  are in conflict with observations [76]. Consequently, it is no longer appropriate to use  ${}^3\text{He}$  as a cosmological probe; instead, one might hope to turn the problem around and constrain stellar astrophysics using the predicted primordial  ${}^3\text{He}$  abundance [77].

### 24.3. Concordance, Dark Matter, and the CMB

We now use the observed light element abundances to test the theory. We first consider standard BBN, which is based on Standard Model physics alone, so  $N_\nu = 3$  and the only free parameter is the baryon-to-photon ratio  $\eta$ . (The implications of BBN for physics beyond the Standard Model will be considered below, Section 24.5). Thus, any abundance measurement determines  $\eta$ , and additional measurements overconstrain the theory and thereby provide a consistency check.

While the  $\eta$  ranges spanned by the boxes in Fig. 24.1 do not all overlap, they are all within a factor  $\sim 2$  of each other. In particular, the lithium abundance corresponds to  $\eta$  values that are inconsistent with that of the (now very precise) D/H abundance as well as the less-constraining  ${}^4\text{He}$  abundance. This discrepancy marks the “lithium problem”. The problem could simply reflect difficulty in determining the primordial lithium abundance; or could hint at a more fundamental omission in the theory. The possibility that lithium reveals new physics is addressed in detail in the next section. If however we exclude the lithium constraint because its inferred abundance may suffer from systematic uncertainties, then D/H and  ${}^4\text{He}$  are in agreement. The concordant  $\eta$  range is essentially that implied by D/H, namely

$$5.8 \leq \eta_{10} \leq 6.6 \text{ (95\% CL)}. \quad (24.5)$$

Despite the lithium problem, the overall concordance remains remarkable: using only well-established microphysics we can extrapolate back to  $t \sim 1$  s to predict light element abundances spanning 9 orders of magnitude, in approximate agreement with observation. This is a major success for the standard cosmology, and inspires confidence in extrapolation back to still earlier times.

This concordance provides a measure of the baryon content:

$$0.021 \leq \Omega_b h^2 \leq 0.024 \text{ (95\% CL)}, \quad (24.6)$$

a result that plays a key role in our understanding of the matter budget of the Universe. First we note that  $\Omega_b \ll 1$ , *i.e.*, baryons cannot close the Universe [78]. Furthermore, the cosmic density of (optically) luminous matter is  $\Omega_{\text{lum}} \simeq 0.0024 h^{-1}$  [79], so that  $\Omega_b \gg \Omega_{\text{lum}}$ : most baryons are optically dark, probably in the form of a diffuse intergalactic medium [80]. Finally, given that  $\Omega_m \sim 0.3$  (see Dark Matter and Cosmological Parameters reviews), we infer that most matter in the Universe is not only dark, but also takes some non-baryonic (more precisely, non-nucleonic) form.

The BBN prediction for the cosmic baryon density can be tested through precision observations of CMB temperature fluctuations (see Cosmic Microwave Background review). One can determine  $\eta$  from the amplitudes of the acoustic peaks in the CMB angular power spectrum [81], making it possible to compare two measures of  $\eta$  using very different physics, at two widely separated epochs. In the standard cosmology, there is no change in  $\eta$  between BBN and CMB decoupling, thus, a comparison of  $\eta_{\text{BBN}}$  and  $\eta_{\text{CMB}}$  is a key test. Agreement would endorse the standard picture, while disagreement could point to new physics during/between the BBN and CMB epochs.

The analysis described in the Cosmic Microwave Background review, based on *Planck* 2015 data, yields  $\Omega_b h^2 = 0.0223 \pm 0.0002$  which corresponds to  $\eta_{10} = 6.09 \pm 0.06$  [55]. This result depends weakly on the primordial helium abundance, and the fiducial *Planck*

analysis uses BBN theory to fix  $Y_p(\eta)$ . As shown in Fig. 24.1, this CMB estimate of the baryon density (narrow vertical band) is remarkably consistent with the BBN range quoted in Eq. (24.6) and thus in very good agreement with the value inferred from recent high-redshift D/H measurements [37] and  ${}^4\text{He}$  determinations; together these observations span diverse environments from redshifts  $z = 1000$  to the present [82].

The CMB damping tail is sensitive to the primordial  ${}^4\text{He}$  abundance, and is independent from both BBN and local  ${}^4\text{He}$  measurements. [54]. The *Planck* 2015 analysis using TT+lowP but not lensing yields  $Y_p = 0.253^{+0.041}_{-0.042}$  [55], *i.e.*, consistent with the H II region helium abundance determination. Moreover, this value is consistent with the Standard ( $N_\nu = 3$ ) BBN prediction for  $Y_p$  at the *Planck*-determined baryon density. This concordance represents a successful CMB-only test of BBN.

The precision determinations of the baryon density using the CMB motivates the use of this value as an input to BBN calculations. Within the context of the Standard Model, BBN then becomes a zero-parameter theory, and the light element abundances are completely determined to within the uncertainties in  $\eta_{\text{CMB}}$  and the BBN theoretical errors. Comparison with the observed abundances then can be used to test the astrophysics of post-BBN light element evolution [83]. Alternatively, one can consider possible physics beyond the Standard Model (*e.g.*, which might change the expansion rate during BBN) and then use all of the abundances to test such models; this is discussed in Section 24.5.

### 24.4. The Lithium Problem

As Fig. 24.1 shows, stellar Li/H measurements are inconsistent with the CMB (and D/H), given the error budgets we have quoted. Recent updates in nuclear cross sections and stellar abundance systematics *increase* the discrepancy to over  $5\sigma$ , depending on the stellar abundance analysis adopted [14].

The question then becomes pressing as to whether this mismatch comes from systematic errors in the observed abundances, and/or uncertainties in stellar astrophysics or nuclear inputs, or whether there might be new physics at work [9]. Nuclear inputs (cross sections) for BBN reactions are constrained by extensive laboratory measurements; to increase  ${}^7\text{Be}$  destruction requires enhancement of otherwise subdominant processes that can be attained by missed resonances in a few reactions such as  ${}^7\text{Be}(d,p)2\alpha$  if the compound nuclear state properties are particularly favorable [84]. However, experimental searches have now closed off these cases [85], making a “nuclear fix” increasingly unlikely.

Another conventional means to solve the lithium problem is by *in situ* destruction over the long lifetimes of the host halo stars. Stellar depletion mechanisms include diffusion, rotationally induced mixing, or pre-main-sequence depletion. These effects certainly occur, but to reduce lithium to the required levels generally requires some *ad hoc* mechanism and fine tuning of the initial stellar parameters [70,67,86]. A putative signature of diffusion has been reported for the globular clusters NGC 6397 and NGC 6752, where the ‘turnoff’ stars exhibit slightly lower (by  $\sim 0.1$  dex) abundances of Fe II, Ti II, Sc II, Ca I and Mg I, than in more evolved stars [69,87]. General features of diffusive models are a dispersion in the Li abundances and a pronounced downturn in the Li abundances at the hot end of the Li plateau. Some extra turbulence needs to be invoked to limit diffusion in the hotter stars and to restore uniform Li abundance along the Spite plateau [86]. In the framework of these models (and also assuming identical initial stellar rotation) depletion by at most a factor  $\sim 1.8$  is conceivable [69,87].

As nuclear and astrophysical solutions to the lithium problem become increasingly constrained (even if difficult to rule out definitively), the possibility of new physics arises. Nucleosynthesis models in which the baryon-to-photon ratio is inhomogeneous can alter abundances for a given  $\eta_{\text{BBN}}$ , but will overproduce  ${}^7\text{Li}$  [88]. Entropy generation by some non-standard process could have decreased  $\eta$  between the BBN era and CMB decoupling, however the lack of spectral distortions in the CMB rules out any significant energy

injection upto a redshift  $z \sim 10^7$  [89]. The most intriguing resolution of the lithium problem thus involves new physics during BBN [7–9].

We summarize the general features of such solutions here, and later consider examples in the context of specific particle physics models. Many proposed solutions introduce perturbations to light-element formation during BBN; while all element abundances may suffer perturbations, the interplay of  ${}^7\text{Li}$  and D is often the most important *i.e.* observations of D often provide the strongest constraints on the allowed perturbations to  ${}^7\text{Li}$ . In this connection it is important to note that the new, very precise determination of D/H [37] will significantly constrain the ability of such models to ameliorate or solve the lithium problem.

A well studied class of models invokes the injection of suprathermal hadronic or electromagnetic particles due to decays of dark matter particles. The effects are complex and depend on the nature of the decaying particles and their branchings and spectra. However, the models that most successfully solve the lithium problem generally feature non-thermal nucleons, which dissociate all light elements. Dissociation of even a small fraction of  ${}^4\text{He}$  introduces a large abundance of free neutrons, which quickly thermalize. The thermal neutrons drive the  ${}^7\text{Be}(n,p){}^7\text{Li}$  conversion of  ${}^7\text{Be}$ . The resulting  ${}^7\text{Li}$  has a lower Coulomb barrier relative to  ${}^7\text{Be}$  and is readily destroyed via  ${}^7\text{Li}(p,\alpha){}^4\text{He}$  [90,91]. But  ${}^4\text{He}$  dissociation also produces D directly and via nonthermal neutron  $n(p,\gamma)d$  reactions; this introduces a tension between Li/H reduction and D/H enhancement.

Another important class of models retains the standard cosmic particle content, but changes their interactions via time variations in the fundamental constants [92]. Here too, the details are model-dependent, but scenarios that solve or alleviate the lithium problem often feature perturbations to the deuteron binding energy. A weaker D binding leads to the D bottleneck being overcome later, so that element formation commences at a lower temperature and lower density. This leads in turn to slower nuclear rates that freeze out earlier. The net result is a *higher* final D/H, due to less efficient processing into  ${}^4\text{He}$ , but also *lower* Li, due to suppressed production via  ${}^3\text{He}(\alpha,\gamma){}^7\text{Be}$ .

The lithium problem remains an unresolved issue in BBN. Nevertheless, the remarkable concordance between the CMB and the D (as well as  ${}^4\text{He}$ ) abundance, remains a non-trivial success, and provides constraints on the early Universe and particle physics.

## 24.5. Beyond the Standard Model

Given the simple physics underlying BBN, it is remarkable that it still provides the most effective test for the cosmological viability of ideas concerning physics beyond the Standard Model. Although baryogenesis and inflation must have occurred at higher temperatures in the early Universe, we do not as yet have ‘standard models’ for these, so BBN still marks the boundary between the established and the speculative in Big Bang cosmology. It might appear possible to push the boundary back to the quark-hadron transition at  $T \sim \Lambda_{\text{QCD}}$ , or electroweak symmetry breaking at  $T \sim 1/\sqrt{G_F}$ ; however, so far no observable relics of these epochs have been identified, either theoretically or observationally. Thus, although the Standard Model provides a precise description of physics up to the Fermi scale, cosmology cannot be traced in detail before the BBN era.

Limits on new physics come mainly from the observational bounds on the  ${}^4\text{He}$  abundance. This is proportional to the  $n/p$  ratio when the weak-interaction rate falls behind the Hubble expansion rate at  $T_{\text{fr}} \sim 1$  MeV. The presence of additional neutrino flavors (or of any other relativistic species) at this time increases  $g_*$ , hence the expansion rate, leading to a larger value of  $T_{\text{fr}}$ ,  $n/p$ , and therefore  $Y_p$  [10,93]. In the Standard Model at  $T = 1$  MeV,  $g_* = 5.5 + \frac{7}{4}N_\nu$ , where  $N_\nu$  is the *effective* number of (nearly) massless neutrino flavors (see Big Bang Cosmology review). The helium curves in Fig. 24.1 were computed taking  $N_\nu = 3$ ; small corrections for non-equilibrium neutrino heating [20] are included in the thermal evolution and lead to an effective  $N_\nu = 3.04$  compared to assuming instantaneous neutrino freezeout (see Big Bang Cosmology review). The computed  ${}^4\text{He}$  abundance scales as  $\Delta Y_p \simeq 0.013 \Delta N_\nu$  [11]. Clearly the central value for  $N_\nu$  from BBN will depend on  $\eta$ , which is independently

determined (with weaker sensitivity to  $N_\nu$ ) by the adopted D or  ${}^7\text{Li}$  abundance. For example, if the best value for the observed primordial  ${}^4\text{He}$  abundance is 0.249, then, for  $\eta_{10} \sim 6$ , the central value for  $N_\nu$  is very close to 3. A maximum likelihood analysis on  $\eta$  and  $N_\nu$  based on the above  ${}^4\text{He}$  and D abundances finds the (correlated) 95% CL ranges to be  $5.6 < \eta_{10} < 6.6$  and  $2.3 < N_\nu < 3.4$  [5]. Identical results are obtained using a simpler method to extract such bounds based on  $\chi^2$  statistics, given a set of input abundances [94].

The CMB power spectrum in the damping tail is independently sensitive to  $N_\nu$  (*e.g.* [95]). The CMB value  $N_\nu^{\text{CMB}}$  probes the cosmic radiation content at (re)combination, so a discrepancy would imply new physics or astrophysics. Indeed, observations by the South Pole Telescope implied  $N_\nu^{\text{CMB}} = 3.85 \pm 0.62$  [96], prompting discussion of ‘dark radiation’ such as sterile neutrinos [97]. However, *Planck* 2015 results give  $N_\nu^{\text{CMB}} = 3.13 \pm 0.31$  when using the BBN  $Y_p(\eta)$ , a result quite consistent with the Standard Model neutrinos [55]. If we *assume* that  $\eta$  did not change between BBN and (re)combination, the constraint can be improved by including the recent D/H and astrophysical  $Y_p$  measurements, which yields  $N_\nu = 2.88 \pm 0.16$  [5].

Just as one can use the measured helium abundance to place limits on  $g_*$  [93], any changes in the strong, weak, electromagnetic, or gravitational coupling constants, arising *e.g.*, from the dynamics of new dimensions, can be similarly constrained [98], as can be any speed-up of the expansion rate in, *e.g.*, scalar-tensor theories of gravity [99].

The limits on  $N_\nu$  can be translated into limits on other types of particles or particle masses that would affect the expansion rate of the Universe during nucleosynthesis. For example, consider ‘sterile’ neutrinos with only right-handed interactions of strength  $G_R < G_F$ . Such particles would decouple at higher temperature than (left-handed) neutrinos, so their number density ( $\propto T^3$ ) relative to neutrinos would be reduced by any subsequent entropy release, *e.g.*, due to annihilations of massive particles that become non-relativistic between the two decoupling temperatures. Thus (relativistic) particles with less than full strength weak interactions contribute less to the energy density than particles that remain in equilibrium up to the time of nucleosynthesis [100]. If we impose  $N_\nu < 4$  as an illustrative constraint, then the three right-handed neutrinos must have a temperature  $3(T_{\nu_R}/T_{\nu_L})^4 < 1$ . Since the temperature of the decoupled  $\nu_R$  is determined by entropy conservation (see Big Bang Cosmology review),  $T_{\nu_R}/T_{\nu_L} = [(43/4)/g_*(T_d)]^{1/3} < 0.76$ , where  $T_d$  is the decoupling temperature of the  $\nu_R$ . This requires  $g_*(T_d) > 24$ , so decoupling must have occurred at  $T_d > 140$  MeV. The decoupling temperature is related to  $G_R$  through  $(G_R/G_F)^2 \sim (T_d/3 \text{ MeV})^{-3}$ , where 3 MeV is the decoupling temperature for  $\nu_L$ s. This yields a limit  $G_R \lesssim 10^{-2} G_F$ . The above argument sets lower limits on the masses of new  $Z'$  gauge bosons to which right-handed neutrinos would be coupled in models of superstrings [101], or extended technicolor [102]. Similarly a Dirac magnetic moment for neutrinos, which would allow the right-handed states to be produced through scattering and thus increase  $g_*$ , can be significantly constrained [103], as can any new interactions for neutrinos that have a similar effect [104]. Right-handed states can be populated directly by helicity-flip scattering if the neutrino mass is large enough, and this property has been used to infer a bound of  $m_{\nu_\tau} \lesssim 1$  MeV taking  $N_\nu < 4$  [105]. If there is mixing between active and sterile neutrinos then the effect on BBN is more complicated [106].

BBN limits on the cosmic expansion rate constrain supersymmetric scenarios in which the neutralino or gravitino are very light, so that they contribute to  $g_*$  [107]. A gravitino in the mass range  $\sim 10^{-4} - 10$  eV will affect the expansion rate of the Universe similarly to a light neutralino (which is however now probably ruled out by collider data, especially the decays of the Higgs-like boson). The net contribution to  $N_\nu$  then ranges between 0.74 and 1.69, depending on the gravitino and slepton masses [108].

The limit on the expansion rate during BBN can also be translated into bounds on the mass/lifetime of non-relativistic particles that decay during BBN. This results in an even faster speed-up rate, and typically also changes the entropy [109]. If the decays include Standard Model particles, the resulting electromagnetic [110–111]



and/or hadronic [112] cascades can strongly perturb the light elements, which leads to even stronger constraints. Such arguments had been applied to rule out an MeV mass for  $\nu_\tau$ , which decays during nucleosynthesis [113].

Decaying-particle arguments have proved very effective in probing supersymmetry. Light-element abundances generally are complementary to accelerator data in constraining SUSY parameter space, with BBN reaching to values kinematically inaccessible to the LHC. Much recent interest has focused on the case in which the next-to-lightest supersymmetric particle is metastable and decays during or after BBN. The constraints on unstable particles discussed above imply stringent bounds on the allowed abundance of such particles [112]; if the metastable particle is charged (*e.g.*, the stau), then it is possible for it to form atom-like electromagnetic bound states with nuclei, and the resulting impact on light elements can be quite complex [114]. Moreover, SUSY decays can destroy  ${}^7\text{Li}$  and/or produce  ${}^6\text{Li}$ , leading to a possible supersymmetric solution to the lithium problems noted above [115] (see [7] for a review).

These arguments impose powerful constraints on supersymmetric inflationary cosmology [111–112], particularly thermal leptogenesis [116]. These can be evaded only if the gravitino is massive enough to decay before BBN, *i.e.*,  $m_{3/2} \gtrsim 50$  TeV [117] (which would be unnatural), or if it is in fact the lightest supersymmetric particle and thus stable [111,118]. Similar constraints apply to moduli – very weakly coupled fields in string theory that obtain an electroweak-scale mass from supersymmetry breaking [119].

Finally, we mention that BBN places powerful constraints on the possibility that there are new large dimensions in nature, perhaps enabling the scale of quantum gravity to be as low as the electroweak scale [120]. Thus, Standard Model fields may be localized on a ‘brane,’ while gravity alone propagates in the ‘bulk.’ It has been further noted that the new dimensions may be non-compact, even infinite [121], and the cosmology of such models has attracted considerable attention. The expansion rate in the early Universe can be significantly modified, so BBN is able to set interesting constraints on such possibilities [122].

## References:

- R.V. Wagoner *et al.*, *Astrophys. J.* **148**, 3 (1967).
- D.N. Schramm and M.S. Turner, *Rev. Mod. Phys.* **70**, 303 (1998).
- G. Steigman, *Ann. Rev. Nucl. and Part. Sci.* **57**, 463 (2007).
- F. Iocco *et al.*, *Phys. Reports* **472**, 1 (2009).
- R.H. Cyburt *et al.*, *Rev. Mod. Phys.* **88**, 015004 (2016).
- S. Sarkar, *Rept. on Prog. in Phys.* **59**, 1493 (1996).
- K. Jedamzik and M. Pospelov, *New J. Phys.* **11**, 105028 (2009).
- M. Pospelov and J. Pradler, *Ann. Rev. Nucl. and Part. Sci.* **60**, 539 (2010).
- B.D. Fields, *Ann. Rev. Nucl. and Part. Sci.* **61**, 47 (2011).
- P.J.E. Peebles, *Phys. Rev. Lett.* **16**, 411 (1966).
- J. Bernstein *et al.*, *Rev. Mod. Phys.* **61**, 25 (1989).
- S. Mukhanov, *Int. J. Theor. Phys.* **143**, 669 (2004).
- R. Esmailzadeh *et al.*, *Astrophys. J.* **378**, 504 (1991).
- R.H. Cyburt *et al.*, *JCAP* **0811**, 012 (2008).
- R.H. Cyburt *et al.*, *New Astron.* **6**, 215 (2001).
- L. Kawano, *FERMILAB-PUB-92/04-A*.
- O. Pisanti *et al.*, *Comp. Phys. Comm.* **178**, 956 (2008).
- A. Arbey, *Comp. Phys. Comm.* **183**, 1822 (2012).
- S. Esposito *et al.*, *Nucl. Phys.* **B568**, 421 (2000).
- S. Dodelson and M.S. Turner, *Phys. Rev.* **D46**, 3372 (1992).
- D. Seckel, [hep-ph/9305311](#);  
R. Lopez and M.S. Turner, *Phys. Rev.* **D59**, 103502 (1999).
- M.S. Smith *et al.*, *Astrophys. J. Supp.* **85**, 219 (1993).
- G. Fiorentini *et al.*, *Phys. Rev.* **D58**, 063506 (1998).
- A. Coc *et al.*, *Astrophys. J.* **744**, 158 (2012).
- K.M. Nollett and S. Burles, *Phys. Rev.* **D61**, 123505 (2000).
- R.H. Cyburt, *Phys. Rev.* **D70**, 023505 (2004).
- P.D. Serpico *et al.*, *JCAP* **12**, 010 (2004).
- R.N. Boyd *et al.*, *Phys. Rev.* **D82**, 105005 (2010).
- R.H. Cyburt and B. Davids, *Phys. Rev.* **C78**, 012 (2008).
- K.M. Nollett *et al.*, *Astrophys. J. Lett.* **552**, L1 (2001).
- K.M. Nollett and G.P. Holder, [arXiv:1112.2683](#).
- R.I. Epstein *et al.*, *Nature* **263**, 198 (1976).
- B.E. Wood *et al.*, *Astrophys. J.* **609**, 838 (2004).
- S. D’Odorico *et al.*, *Astron. & Astrophys.* **368**, L21 (2001).
- D. Romano *et al.*, *MNRAS* **369**, 295 (2006).
- M. Pettini and D. Bowen, *Astrophys. J.* **560**, 41 (2001).
- R. Cooke *et al.*, *Astrophys. J.* **781**, 31 (2014).
- S.A. Levshakov *et al.*, *Astrophys. J.* **565**, 696 (2002).
- M. Fumagalli *et al.*, *Science* **334**, 1245 (2011).
- R. Srianand *et al.*, *MNRAS* **405**, 1888 (2010).
- P. Noterdaeme *et al.*, *Astron. & Astrophys.* **542**, L33 (2012).
- M. Pettini and R. Cooke, *MNRAS* **425**, 2477 (2012).
- S. Riemer-Sørensen *et al.*, *MNRAS* **447**, 2925 (2015).
- J.L. Linsky *et al.*, *Astrophys. J.* **647**, 1106 (2006).
- B.D. Savage *et al.*, *Astrophys. J.* **659**, 1222 (2007).
- Y.I. Izotov *et al.*, *Astrophys. J.* **527**, 757 (1999).
- K.A. Olive and E. Skillman, *Astrophys. J.* **617**, 29 (2004).
- M. Peimbert *et al.*, *Astrophys. J.* **667**, 636 (2007).
- Y.I. Izotov *et al.*, *Astrophys. J.* **662**, 15 (2007).
- E. Aver *et al.*, *JCAP* **04**, 004 (2012).
- E. Aver *et al.*, *JCAP* **07**, 011 (2015).
- Y.I. Izotov *et al.*, *MNRAS* **445**, 778 (2014).
- Y.I. Izotov *et al.*, *Astron. & Astrophys.* **558**, A57 (2013).
- R. Trotta and S.H. Hansen, *Phys. Rev.* **D69**, 023509 (2004).
- Planck Collab. 2015 Results XIII, P.A.R. Ade *et al.*, [arXiv:1502.01589](#).
- N. Christlieb *et al.*, *Nature* **419**, 904 (2002).
- M. Spite and F. Spite, *Nature* **297**, 483 (1982).
- E. Vangioni-Flam *et al.*, *New Astron.* **4**, 245 (1999).
- S.G. Ryan *et al.*, *Astrophys. J. Lett.* **530**, L57 (2000).
- P. Bonifacio and P. Molaro, *MNRAS* **285**, 847 (1997).
- P. Bonifacio *et al.*, *Astron. & Astrophys.* **390**, 91 (2002).
- J. Melendez *et al.*, *Astron. & Astrophys.* **515**, L3 (2010).
- P. Bonifacio *et al.*, *Astron. & Astrophys.* **462**, 851 (2007).
- W. Aoki, *Astrophys. J.* **698**, 1803 (2009);  
A. Hosford *et al.*, *Astron. & Astrophys.* **493**, 601 (2009).
- L. Sbordonc *et al.*, *Astron. & Astrophys.* **522**, A26 (2010).
- E. Caffau *et al.*, *Nature* **477**, 67 (2011).
- P. Molaro *et al.*, *Memorie della Soc. Astronomica Italiana Supp.* **22**, 233 (2012).
- M.H. Pinsonneault *et al.*, *Astrophys. J.* **574**, 389 (2002).
- A.J. Korn *et al.*, *Nature* **442**, 657 (2006).
- X. Fu *et al.*, *MNRAS* **452**, 3256 (2015).
- M. Asplund *et al.*, *Astrophys. J.* **644**, 229 (2006).
- K. Lind *et al.*, *Astron. & Astrophys.* **554**, 96 (2013).
- R. Cayrel *et al.*, *Astron. & Astrophys.* **473**, L37 (2007).
- M. Steffen *et al.*, *Memorie della Soc. Astronomica Italiana Supp.* **22**, 152 (2012).
- T.M. Bania *et al.*, *Nature* **415**, 54 (2002).
- K.A. Olive *et al.*, *Astrophys. J.* **479**, 752 (1997).
- E. Vangioni-Flam *et al.*, *Astrophys. J.* **585**, 611 (2003).
- H. Reeves *et al.*, *Astrophys. J.* **179**, 909 (1973).
- M. Fukugita and P.J.E. Peebles, *Astrophys. J.* **616**, 643 (2004).
- R. Cen and J.P. Ostriker, *Astrophys. J.* **514**, 1 (1999).
- G. Jungman *et al.*, *Phys. Rev.* **D54**, 1332 (1996).
- A. Coc *et al.*, [arXiv:1307.6955](#).
- R.H. Cyburt *et al.*, *Phys. Lett.* **B567**, 227 (2003).
- R.H. Cyburt and M. Pospelov, *Int. J. Mod. Phys. E* **21**, 1250004 (2012);  
R.N. Boyd *et al.*, *Phys. Rev.* **D82**, 105005 (2010);  
N. Chakraborty *et al.*, *Phys. Rev.* **D83**, 063006 (2011);  
C. Broggini *et al.*, *JCAP* **06**, 030 (2012).
- P.D. O’Malley *et al.*, *Phys. Rev.* **C84**, 042801 (2011);  
Hammache, F., *et al.*, *Phys. Rev. C* **88**, 062802 (2013);  
Paris, M. *et al.*, *Nuclear Data Sheets*, **120**, 184 (2014).
- O. Richard *et al.*, *Astrophys. J.* **619**, 538 (2005).
- P. Gruyters *et al.*, *Astron. Astrophys.* **555**, 31 (2013).
- K. Jedamzik and J.B. Rehm, *Phys. Rev.* **D64**, 023510 (2001).
- D.J. Fixsen *et al.*, *Astrophys. J.* **473**, 576 (1996).
- K. Jedamzik, *Phys. Rev.* **D70**, 063524 (2004).
- M. Kawasaki *et al.*, *Phys. Rev.* **D71**, 083502 (2005).



92. J.D. Barrow, Phys. Rev. **D35**, 1805 (1987);  
B.A. Campbell and K.A. Olive, Phys. Lett. **B345**, 429 (1995);  
L. Bergström, Phys. Rev. **D60**, 045005 (1999);  
V.V. Flambaum and E.V. Shuryak, Phys. Rev. **D65**, 103503 (2002);  
A. Coc *et al.*, Phys. Rev. **D76**, 023511 (2007);  
J.C. Berengut *et al.*, Phys. Rev. **D87**, 085018 (2013).
93. G. Steigman *et al.*, Phys. Lett. **B66**, 202 (1977).
94. E. Lisi *et al.*, Phys. Rev. **D59**, 123520 (1999).
95. Z. Hou *et al.*, Phys. Rev. **D87**, 083008 (2013).
96. R. Keisler *et al.*, Astrophys. J. **743**, 28 (2011).
97. J. Hamann *et al.*, Phys. Rev. Lett. **105**, 181301 (2010).
98. E.W. Kolb *et al.*, Phys. Rev. **D33**, 869 (1986);  
F.S. Accetta *et al.*, Phys. Lett. **B248**, 146 (1990);  
B.A. Campbell and K.A. Olive, Phys. Lett. **B345**, 429 (1995);  
K.M. Nollett and R. Lopez, Phys. Rev. **D66**, 063507 (2002);  
C. Bambi *et al.*, Phys. Rev. **D71**, 123524 (2005).
99. A. Coc *et al.*, Phys. Rev. **D73**, 083525 (2006).
100. K.A. Olive *et al.*, Nucl. Phys. **B180**, 497 (1981).
101. J. Ellis *et al.*, Phys. Lett. **B167**, 457 (1986).
102. L.M. Krauss *et al.*, Phys. Rev. Lett. **71**, 823 (1993).
103. J.A. Morgan, Phys. Lett. **B102**, 247 (1981).
104. E.W. Kolb *et al.*, Phys. Rev. **D34**, 2197 (1986);  
J.A. Grifols and E. Massó, Mod. Phys. Lett. **A2**, 205 (1987);  
K.S. Babu *et al.*, Phys. Rev. Lett. **67**, 545 (1991).
105. A.D. Dolgov *et al.*, Nucl. Phys. **B524**, 621 (1998).
106. K. Enqvist *et al.*, Nucl. Phys. **B373**, 498 (1992);  
A.D. Dolgov, Phys. Reports **370**, 333 (2002).
107. J.A. Grifols *et al.*, Phys. Lett. **B400**, 124 (1997).
108. H. Dreiner *et al.*, Phys. Rev. **D85**, 065027 (2012).
109. K. Sato and M. Kobayashi, Prog. Theor. Phys. **58**, 1775 (1977);  
D.A. Dicus *et al.*, Phys. Rev. **D17**, 1529 (1978);  
R.J. Scherrer and M.S. Turner, Astrophys. J. **331**, 19 (1988).
110. D. Lindley, MNRAS **188**, 15 (1979); Astrophys. J. **294**, 1 (1985).
111. J. Ellis *et al.*, Nucl. Phys. **B259**, 175 (1985);  
J. Ellis *et al.*, Nucl. Phys. **B373**, 399 (1992);  
R.H. Cyburt *et al.*, Phys. Rev. **D67**, 103521 (2003).
112. M.H. Reno and D. Seckel, Phys. Rev. **D37**, 3441 (1988);  
S. Dimopoulos *et al.*, Nucl. Phys. **B311**, 699 (1989);  
K. Kohri *et al.*, Phys. Rev. **D71**, 083502 (2005).
113. S. Sarkar and A.M. Cooper, Phys. Lett. **B148**, 347 (1984).
114. M. Pospelov *et al.*, Phys. Rev. Lett. **98**, 231301 (2007);  
M. Kawasaki *et al.*, Phys. Lett. **B649**, 436 (2007);  
R.H. Cyburt *et al.*, JCAP **05**, 014 (2013).
115. K. Jedamzik *et al.*, JCAP **07**, 007 (2006).
116. S. Davidson *et al.*, Phys. Rev. **466**, 105 (2008).
117. S. Weinberg, Phys. Rev. Lett. **48**, 1303 (1979).
118. M. Bolz *et al.*, Nucl. Phys. **B606**, 518 (2001).
119. G. Coughlan *et al.*, Phys. Lett. **B131**, 59 (1983).
120. N. Arkani-Hamed *et al.*, Phys. Rev. **D59**, 086004 (1999).
121. L. Randall and R. Sundrum, Phys. Rev. Lett. **83**, 3370 (1999).
122. J.M. Cline *et al.*, Phys. Rev. Lett. **83**, 4245 (1999);  
P. Binetruy *et al.*, Phys. Lett. **B477**, 285 (2000).

## 25. THE COSMOLOGICAL PARAMETERS

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### 25.1. Parametrizing the Universe

Rapid advances in observational cosmology have led to the establishment of a precision cosmological model, with many of the key cosmological parameters determined to one or two significant figure accuracy. Particularly prominent are measurements of cosmic microwave background (CMB) anisotropies, with the highest precision observations being those of the *Planck* Satellite [1,2] which supersede the iconic *WMAP* results [3,4]. However the most accurate model of the Universe requires consideration of a range of observations, with complementary probes providing consistency checks, lifting parameter degeneracies, and enabling the strongest constraints to be placed.

The term ‘cosmological parameters’ is forever increasing in its scope, and nowadays often includes the parameterization of some functions, as well as simple numbers describing properties of the Universe. The original usage referred to the parameters describing the global dynamics of the Universe, such as its expansion rate and curvature. Also now of great interest is how the matter budget of the Universe is built up from its constituents: baryons, photons, neutrinos, dark matter, and dark energy. We need to describe the nature of perturbations in the Universe, through global statistical descriptors such as the matter and radiation power spectra. There may also be parameters describing the physical state of the Universe, such as the ionization fraction as a function of time during the era since recombination. Typical comparisons of cosmological models with observational data now feature between five and ten parameters.

#### 25.1.1. The global description of the Universe :

Ordinarily, the Universe is taken to be a perturbed Robertson–Walker space-time with dynamics governed by Einstein’s equations. This is described in detail in the Big-Bang Cosmology chapter in this volume. Using the density parameters  $\Omega_i$  for the various matter species and  $\Omega_\Lambda$  for the cosmological constant, the Friedmann equation can be written

$$\sum_i \Omega_i + \Omega_\Lambda - 1 = \frac{k}{R^2 H^2}, \quad (25.1)$$

where the sum is over all the different species of material in the Universe. This equation applies at any epoch, but later in this article we will use the symbols  $\Omega_i$  and  $\Omega_\Lambda$  to refer to the present-epoch values.

The complete present-epoch state of the homogeneous Universe can be described by giving the current-epoch values of all the density parameters and the Hubble constant  $h$  (the present-day Hubble parameter being written  $H_0 = 100h \text{ km s}^{-1} \text{ Mpc}^{-1}$ ). A typical collection would be baryons  $\Omega_b$ , photons  $\Omega_\gamma$ , neutrinos  $\Omega_\nu$ , and cold dark matter  $\Omega_c$  (given charge neutrality, the electron density is guaranteed to be too small to be worth considering separately and is effectively included with the baryons). The spatial curvature can then be determined from the other parameters using Eq. (25.1). The total present matter density  $\Omega_m = \Omega_c + \Omega_b$  is sometimes used in place of the cold dark matter density  $\Omega_c$ .

These parameters also allow us to track the history of the Universe, at least back until an epoch where interactions allow interchanges between the densities of the different species; this is believed to have last happened at neutrino decoupling, shortly before Big Bang Nucleosynthesis (BBN). To probe further back into the Universe’s history requires assumptions about particle interactions, and perhaps about the nature of physical laws themselves.

The standard neutrino sector has three flavors. For neutrinos of mass in the range  $5 \times 10^{-4} \text{ eV}$  to  $1 \text{ MeV}$ , the density parameter in neutrinos is predicted to be

$$\Omega_\nu h^2 = \frac{\sum m_\nu}{93 \text{ eV}}, \quad (25.2)$$

where the sum is over all families with mass in that range (higher masses need a more sophisticated calculation). We use units with  $c = 1$  throughout. Results on atmospheric and Solar neutrino oscillations [5] imply non-zero mass-squared differences between the three neutrino

flavors. These oscillation experiments cannot tell us the absolute neutrino masses, but within the simple assumption of a mass hierarchy suggest a lower limit of approximately  $0.06 \text{ eV}$  for the sum of the neutrino masses (see the Neutrino section).

Even a mass this small has a potentially observable effect on the formation of structure, as neutrino free-streaming damps the growth of perturbations. Analyses commonly now either assume a neutrino mass sum fixed at this lower limit, or allow the neutrino mass sum as a variable parameter. To date there is no decisive evidence of any effects from either neutrino masses or an otherwise non-standard neutrino sector, and observations impose quite stringent limits, which we summarize in Section 25.3.4. However, we note that the inclusion of the neutrino mass sum as a free parameter can affect the derived values of other cosmological parameters.

#### 25.1.2. Inflation and perturbations :

A complete model of the Universe should include a description of deviations from homogeneity, at least in a statistical way. Indeed, some of the most powerful probes of the parameters described above come from the evolution of perturbations, so their study is naturally intertwined with the determination of cosmological parameters.

There are many different notations used to describe the perturbations, both in terms of the quantity used to describe the perturbations and the definition of the statistical measure. We use the dimensionless power spectrum  $\Delta^2$  as defined in the Big Bang Cosmology chapter (also denoted  $\mathcal{P}$  in some of the literature). If the perturbations obey Gaussian statistics, the power spectrum provides a complete description of their properties.

From a theoretical perspective, a useful quantity to describe the perturbations is the curvature perturbation  $\mathcal{R}$ , which measures the spatial curvature of a comoving slicing of the space-time. A simple case is the Harrison–Zeldovich spectrum, which corresponds to a constant  $\Delta_{\mathcal{R}}^2$ . More generally, one can approximate the spectrum by a power-law, writing

$$\Delta_{\mathcal{R}}^2(k) = \Delta_{\mathcal{R}}^2(k_*) \left[ \frac{k}{k_*} \right]^{n_s - 1}, \quad (25.3)$$

where  $n_s$  is known as the spectral index, always defined so that  $n_s = 1$  for the Harrison–Zeldovich spectrum, and  $k_*$  is an arbitrarily chosen scale. The initial spectrum, defined at some early epoch of the Universe’s history, is usually taken to have a simple form such as this power law, and we will see that observations require  $n_s$  close to one. Subsequent evolution will modify the spectrum from its initial form.

The simplest mechanism for generating the observed perturbations is the inflationary cosmology, which posits a period of accelerated expansion in the Universe’s early stages [6,7]. It is a useful working hypothesis that this is the sole mechanism for generating perturbations, and it may further be assumed to be the simplest class of inflationary model, where the dynamics are equivalent to that of a single scalar field  $\phi$  with canonical kinetic energy slowly rolling on a potential  $V(\phi)$ . One may seek to verify that this simple picture can match observations and to determine the properties of  $V(\phi)$  from the observational data. Alternatively, more complicated models, perhaps motivated by contemporary fundamental physics ideas, may be tested on a model-by-model basis (see more in the Inflation chapter in this volume).

Inflation generates perturbations through the amplification of quantum fluctuations, which are stretched to astrophysical scales by the rapid expansion. The simplest models generate two types, density perturbations that come from fluctuations in the scalar field and its corresponding scalar metric perturbation, and gravitational waves that are tensor metric fluctuations. The former experience gravitational instability and lead to structure formation, while the latter can influence the CMB anisotropies. Defining slow-roll parameters, with primes indicating derivatives with respect to the scalar field, as

$$\epsilon = \frac{m_{\text{Pl}}^2}{16\pi} \left( \frac{V'}{V} \right)^2, \quad \eta = \frac{m_{\text{Pl}}^2}{8\pi} \frac{V''}{V}, \quad (25.4)$$

which should satisfy  $\epsilon, |\eta| \ll 1$ , the spectra can be computed using the slow-roll approximation as

$$\Delta_{\mathcal{R}}^2(k) \simeq \frac{8}{3m_{\text{Pl}}^4} \frac{V}{\epsilon} \bigg|_{k=aH}, \quad \Delta_{\text{t}}^2(k) \simeq \frac{128}{3m_{\text{Pl}}^4} V \bigg|_{k=aH}. \quad (25.5)$$

In each case, the expressions on the right-hand side are to be evaluated when the scale  $k$  is equal to the Hubble radius during inflation. The symbol ‘ $\simeq$ ’ here indicates use of the slow-roll approximation, which is expected to be accurate to a few percent or better.

From these expressions, we can compute the spectral indices [8]:

$$n_s \simeq 1 - 6\epsilon + 2\eta \quad ; \quad n_t \simeq -2\epsilon. \quad (25.6)$$

Another useful quantity is the ratio of the two spectra, defined by

$$r \equiv \frac{\Delta_{\mathcal{R}}^2(k_*)}{\Delta_{\mathcal{R}}^2(k_*)}. \quad (25.7)$$

We have

$$r \simeq 16\epsilon \simeq -8n_t, \quad (25.8)$$

which is known as the consistency equation.

One could consider corrections to the power-law approximation, which we discuss later. However, for now we make the working assumption that the spectra can be approximated by such power laws. The consistency equation shows that  $r$  and  $n_t$  are not independent parameters, and so the simplest inflation models give initial conditions described by three parameters, usually taken as  $\Delta_{\mathcal{R}}^2$ ,  $n_s$ , and  $r$ , all to be evaluated at some scale  $k_*$ , usually the ‘statistical center’ of the range explored by the data. Alternatively, one could use the parametrization  $V$ ,  $\epsilon$ , and  $\eta$ , all evaluated at a point on the putative inflationary potential.

After the perturbations are created in the early Universe, they undergo a complex evolution up until the time they are observed in the present Universe. When the perturbations are small, this can be accurately followed using a linear theory numerical code such as CAMB or CLASS [9]. This works right up to the present for the CMB, but for density perturbations on small scales non-linear evolution is important and can be addressed by a variety of semi-analytical and numerical techniques. However the analysis is made, the outcome of the evolution is in principle determined by the cosmological model and by the parameters describing the initial perturbations, and hence can be used to determine them.

Of particular interest are CMB anisotropies. Both the total intensity and two independent polarization modes are predicted to have anisotropies. These can be described by the radiation angular power spectra  $C_\ell$  as defined in the CMB article in this volume, and again provide a complete description if the density perturbations are Gaussian.

### 25.1.3. The standard cosmological model :

We now have most of the ingredients in place to describe the cosmological model. Beyond those of the previous subsections, we need a measure of the ionization state of the Universe. The Universe is known to be highly ionized at low redshifts (otherwise radiation from distant quasars would be heavily absorbed in the ultra-violet), and the ionized electrons can scatter microwave photons, altering the pattern of observed anisotropies. The most convenient parameter to describe this is the optical depth to scattering  $\tau$  (*i.e.*, the probability that a given photon scatters once); in the approximation of instantaneous and complete reionization, this could equivalently be described by the redshift of reionization  $z_{\text{ion}}$ .

As described in Sec. 25.4, models based on these parameters are able to give a good fit to the complete set of high-quality data available at present, and indeed some simplification is possible. Observations are consistent with spatial flatness, and the inflation models so far described automatically generate negligible spatial curvature, so we can set  $k = 0$ ; the density parameters then must sum to unity, and so one of them can be eliminated. The neutrino energy density is often not taken as an independent parameter. Provided that the neutrino sector has the standard interactions, the neutrino energy density, while relativistic, can be related to the photon density using thermal physics arguments, and a minimal assumption takes the neutrino mass sum to be that of the lowest mass solution to the neutrino oscillation constraints, namely 0.06 eV. In addition, there is no observational evidence for the existence of tensor perturbations (though the upper limits are fairly weak), and so  $r$  could be set to zero. This leaves seven

parameters, which is the smallest set that can usefully be compared to the present cosmological data set. This model is referred to by various names, including  $\Lambda$ CDM, the concordance cosmology, and the standard cosmological model.

Of these parameters, only  $\Omega_\gamma$  is accurately measured directly. The radiation density is dominated by the energy in the CMB, and the COBE satellite FIRAS experiment determined its temperature to be  $T = 2.7255 \pm 0.0006$  K [10], <sup>‡</sup> corresponding to  $\Omega_\gamma = 2.47 \times 10^{-5} h^{-2}$ . It typically need not be varied in fitting other data. Hence the minimum number of cosmological parameters varied in fits to data is six, though as described below there may additionally be many ‘nuisance’ parameters necessary to describe astrophysical processes influencing the data.

In addition to this minimal set, there is a range of other parameters that might prove important in future as the data-sets further improve, but for which there is so far no direct evidence, allowing them to be set to a specific value for now. We discuss various speculative options in the next section. For completeness at this point, we mention one other interesting parameter, the helium fraction, which is a non-zero parameter that can affect the CMB anisotropies at a subtle level. It is usually fixed in microwave anisotropy studies, but the data are approaching a level where allowing its variation may become mandatory.

Most attention to date has been on parameter estimation, where a set of parameters is chosen by hand and the aim is to constrain them. Interest has been growing towards the higher-level inference problem of model selection, which compares different choices of parameter sets. Bayesian inference offers an attractive framework for cosmological model selection, setting a tension between model predictiveness and ability to fit the data [11].

#### 25.1.4. Derived parameters :

The parameter list of the previous subsection is sufficient to give a complete description of cosmological models that agree with observational data. However, it is not a unique parameterization, and one could instead use parameters derived from that basic set. Parameters that can be obtained from the set given above include the age of the Universe, the present horizon distance, the present neutrino background temperature, the epoch of matter–radiation equality, the epochs of recombination and decoupling, the epoch of transition to an accelerating Universe, the baryon-to-photon ratio, and the baryon to dark matter density ratio. In addition, the physical densities of the matter components,  $\Omega_i h^2$ , are often more useful than the density parameters. The density perturbation amplitude can be specified in many different ways other than the large-scale primordial amplitude, for instance, in terms of its effect on the CMB, or by specifying a short-scale quantity, a common choice being the present linear-theory mass dispersion on a scale of  $8 h^{-1} \text{Mpc}$ , known as  $\sigma_8$ .

Different types of observation are sensitive to different subsets of the full cosmological parameter set, and some are more naturally interpreted in terms of some of the derived parameters of this subsection than on the original base parameter set. In particular, most types of observation feature degeneracies whereby they are unable to separate the effects of simultaneously varying specific combinations of several of the base parameters.

### 25.2. Extensions to the standard model

At present, there is no positive evidence in favor of extensions of the standard model. These are becoming increasingly constrained by the data, though there always remains the possibility of trace effects at a level below present observational capability.

<sup>‡</sup> Unless stated otherwise, all quoted uncertainties in this article are one-sigma/68% confidence and all upper limits are 95% confidence. Cosmological parameters sometimes have significantly non-Gaussian uncertainties. Throughout we have rounded central values, and especially uncertainties, from original sources, in cases where they appear to be given to excessive precision.

### 25.2.1. More general perturbations :

The standard cosmology assumes adiabatic, Gaussian perturbations. Adiabaticity means that all types of material in the Universe share a common perturbation, so that if the space-time is foliated by constant-density hypersurfaces, then all fluids and fields are homogeneous on those slices, with the perturbations completely described by the variation of the spatial curvature of the slices. Gaussianity means that the initial perturbations obey Gaussian statistics, with the amplitudes of waves of different wavenumbers being randomly drawn from a Gaussian distribution of width given by the power spectrum. Note that gravitational instability generates non-Gaussianity; in this context, Gaussianity refers to a property of the initial perturbations, before they evolve.

The simplest inflation models, based on one dynamical field, predict adiabatic perturbations and a level of non-Gaussianity that is too small to be detected by any experiment so far conceived. For present data, the primordial spectra are usually assumed to be power laws.

#### 25.2.1.1. Non-power-law spectra:

For typical inflation models, it is an approximation to take the spectra as power laws, albeit usually a good one. As data quality improves, one might expect this approximation to come under pressure, requiring a more accurate description of the initial spectra, particularly for the density perturbations. In general, one can expand  $\ln \Delta_{\mathcal{R}}^2$  as

$$\ln \Delta_{\mathcal{R}}^2(k) = \ln \Delta_{\mathcal{R}}^2(k_*) + (n_{s,*} - 1) \ln \frac{k}{k_*} + \frac{1}{2} \frac{dn_s}{d \ln k} \Big|_* \ln^2 \frac{k}{k_*} + \dots, \quad (25.9)$$

where the coefficients are all evaluated at some scale  $k_*$ . The term  $dn_s/d \ln k|_*$  is often called the running of the spectral index [12]. Once non-power-law spectra are allowed, it is necessary to specify the scale  $k_*$  at which the spectral index is defined.

#### 25.2.1.2. Isocurvature perturbations:

An isocurvature perturbation is one that leaves the total density unperturbed, while perturbing the relative amounts of different materials. If the Universe contains  $N$  fluids, there is one growing adiabatic mode and  $N - 1$  growing isocurvature modes (for reviews see Ref. 7 and Ref. 13). These can be excited, for example, in inflationary models where there are two or more fields that acquire dynamically-important perturbations. If one field decays to form normal matter, while the second survives to become the dark matter, this will generate a cold dark matter isocurvature perturbation.

In general, there are also correlations between the different modes, and so the full set of perturbations is described by a matrix giving the spectra and their correlations. Constraining such a general construct is challenging, though constraints on individual modes are beginning to become meaningful, with no evidence that any other than the adiabatic mode must be non-zero.

#### 25.2.1.3. Seeded perturbations:

An alternative to laying down perturbations at very early epochs is that they are seeded throughout cosmic history, for instance by topological defects such as cosmic strings. It has long been excluded that these are the sole original of structure, but they could contribute part of the perturbation signal, current limits being just a few percent [14]. In particular, cosmic defects formed in a phase transition ending inflation is a plausible scenario for such a contribution.

#### 25.2.1.4. Non-Gaussianity:

Multi-field inflation models can also generate primordial non-Gaussianity (reviewed, *e.g.*, in Ref. 7). The extra fields can either be in the same sector of the underlying theory as the inflaton, or completely separate, an interesting example of the latter being the curvaton model [15]. Current upper limits on non-Gaussianity are becoming stringent, but there remains strong motivation to push down those limits and perhaps reveal trace non-Gaussianity in the data. If non-Gaussianity is observed, its nature may favor an inflationary origin, or a different one such as topological defects.

### 25.2.2. Dark matter properties :

Dark matter properties are discussed in the Dark Matter chapter in this volume. The simplest assumption concerning the dark matter is that it has no significant interactions with other matter, and that its particles have a negligible velocity as far as structure formation is concerned. Such dark matter is described as ‘cold,’ and candidates include the lightest supersymmetric particle, the axion, and primordial black holes. As far as astrophysicists are concerned, a complete specification of the relevant cold dark matter properties is given by the density parameter  $\Omega_c$ , though those seeking to detect it directly are as interested in its interaction properties.

Cold dark matter is the standard assumption and gives an excellent fit to observations, except possibly on the shortest scales where there remains some controversy concerning the structure of dwarf galaxies and possible substructure in galaxy halos. It has long been excluded for all the dark matter to have a large velocity dispersion, so-called ‘hot’ dark matter, as it does not permit galaxies to form; for thermal relics the mass must be above about 1 keV to satisfy this constraint, though relics produced non-thermally, such as the axion, need not obey this limit. However, in future further parameters might need to be introduced to describe dark matter properties relevant to astrophysical observations. Suggestions that have been made include a modest velocity dispersion (warm dark matter) and dark matter self-interactions. There remains the possibility that the dark matter is comprised of two separate components, *e.g.*, a cold one and a hot one, an example being if massive neutrinos have a non-negligible effect.

#### 25.2.3. Relativistic species :

The number of relativistic species in the young Universe (omitting photons) is denoted  $N_{\text{eff}}$ . In the standard cosmological model only the three neutrino species contribute, and its baseline value is assumed fixed at 3.046 (the small shift from 3 is because of a slight predicted deviation from a thermal distribution [16]). However other species could contribute, for example an extra neutrino, possibly of sterile type, or massless Goldstone bosons or other scalars. It is hence interesting to study the effect of allowing this parameter to vary, and indeed although 3.046 is consistent with the data, most analyses currently suggest a somewhat higher value (*e.g.*, Ref. 17).

#### 25.2.4. Dark energy :

While the standard cosmological model given above features a cosmological constant, in order to explain observations indicating that the Universe is presently accelerating, further possibilities exist under the general headings of ‘dark energy’ and ‘modified gravity’. These topics are described in detail in the Dark Energy chapter in this volume. This article focuses on the case of the cosmological constant, as this simple model is a good match to existing data. We note that more general treatments of dark energy/modified gravity will lead to weaker constraints on other parameters.

#### 25.2.5. Complex ionization history :

The full ionization history of the Universe is given by the ionization fraction as a function of redshift  $z$ . The simplest scenario takes the ionization to have the small residual value left after recombination up to some redshift  $z_{\text{ion}}$ , at which point the Universe instantaneously reionizes completely. Then there is a one-to-one correspondence between  $\tau$  and  $z_{\text{ion}}$  (that relation, however, also depending on other cosmological parameters). An accurate treatment of this process will track separate histories for hydrogen and helium. While currently rapid ionization appears to be a good approximation, as data improve a more complex ionization history may need to be considered.

#### 25.2.6. Varying ‘constants’ :

Variation of the fundamental constants of Nature over cosmological times is another possible enhancement of the standard cosmology. There is a long history of study of variation of the gravitational constant  $G_N$ , and more recently attention has been drawn to the possibility of small fractional variations in the fine-structure constant. There is presently no observational evidence for the former, which is tightly constrained by a variety of measurements. Evidence for the latter has been claimed from studies of spectral line shifts in quasar spectra at redshift  $z \approx 2$  [18], but this is presently controversial and

in need of further observational study. See constraints from the CMB in the Planck 2015 papers.

### 25.2.7. Cosmic topology :

The usual hypothesis is that the Universe has the simplest topology consistent with its geometry, for example that a flat Universe extends forever. Observations cannot tell us whether that is true, but they can test the possibility of a non-trivial topology on scales up to roughly the present Hubble scale. Extra parameters would be needed to specify both the type and scale of the topology, for example, a cuboidal topology would need specification of the three principal axis lengths. At present, there is no evidence for non-trivial cosmic topology [19].

## 25.3. Probes

The goal of the observational cosmologist is to utilize astronomical information to derive cosmological parameters. The transformation from the observables to the parameters usually involves many assumptions about the nature of the objects, as well as of the dark sector. Below we outline the physical processes involved in each of the major probes, and the main recent results. The first two subsections concern probes of the homogeneous Universe, while the remainder consider constraints from perturbations.

In addition to statistical uncertainties we note three sources of systematic uncertainties that will apply to the cosmological parameters of interest: (i) due to the assumptions on the cosmological model and its priors (*i.e.*, the number of assumed cosmological parameters and their allowed range); (ii) due to the uncertainty in the astrophysics of the objects (*e.g.*, light curve fitting for supernovae or the mass–temperature relation of galaxy clusters); and (iii) due to instrumental and observational limitations (*e.g.*, the effect of ‘seeing’ on weak gravitational lensing measurements, or beam shape on CMB anisotropy measurements).

These systematics, the last two of which appear as ‘nuisance parameters’, pose a challenging problem to the statistical analysis. We attempt to fit the whole Universe with 6 to 12 parameters, but we might need to include hundreds of nuisance parameters, some of them highly correlated with the cosmological parameters of interest (for example time-dependent galaxy biasing could mimic the growth of mass fluctuations). Fortunately, there is some astrophysical prior knowledge on these effects, and a small number of physically-motivated free parameters would ideally be preferred in the cosmological parameter analysis.

### 25.3.1. Direct measures of the Hubble constant :

In 1929, Edwin Hubble discovered the law of expansion of the Universe by measuring distances to nearby galaxies. The slope of the relation between the distance and recession velocity is defined to be the Hubble constant,  $H_0$ . Astronomers argued for decades about the systematic uncertainties in various methods and derived values over the wide range  $40 \text{ km s}^{-1} \text{ Mpc}^{-1} \lesssim H_0 \lesssim 100 \text{ km s}^{-1} \text{ Mpc}^{-1}$ .

One of the most reliable results on the Hubble constant comes from the Hubble Space Telescope Key Project [20]. This study used the empirical period–luminosity relation for Cepheid variable stars, and calibrated a number of secondary distance indicators—Type Ia Supernovae (SNe Ia), the Tully–Fisher relation, surface-brightness fluctuations, and Type II Supernovae.

The most recent derivation based on this approach utilizes the maser-based distance to NGC4258 to re-calibrate its Cepheid distance scale [21] to obtain  $H_0 = 72.0 \pm 3.0 \text{ km s}^{-1} \text{ Mpc}^{-1}$  [22]. The major sources of uncertainty in this result are due to the heavy element abundance of the Cepheids and the distance to the fiducial nearby galaxy, the Large Magellanic Cloud, relative to which all Cepheid distances are measured.

The indirect determination of  $H_0$  by the *Planck* Collaboration [2] found a lower value,  $H_0 = 67.8 \pm 0.9 \text{ km s}^{-1} \text{ Mpc}^{-1}$ . As discussed in that paper, there is strong degeneracy of  $H_0$  with other parameters, *e.g.*,  $\Omega_m$  and the neutrino mass. The tension between the  $H_0$  from *Planck* and the traditional cosmic distance-ladder methods is under investigation.

### 25.3.2. Supernovae as cosmological probes :

Empirically, the peak luminosity of SNe Ia can be used as an efficient distance indicator (*e.g.*, Ref. 23), thus allowing cosmology to be constrained via the distance–redshift relation. The favorite theoretical explanation for SNe Ia is the thermonuclear disruption of carbon–oxygen white dwarfs. Although not perfect ‘standard candles’, it has been demonstrated that by correcting for a relation between the light-curve shape, color, and luminosity at maximum brightness, the dispersion of the measured luminosities can be greatly reduced. There are several possible systematic effects that may affect the accuracy of the use of SNe Ia as distance indicators, *e.g.*, evolution with redshift and interstellar extinction in the host galaxy and in the Milky Way.

Two major studies, the Supernova Cosmology Project and the High- $z$  Supernova Search Team, found evidence for an accelerating Universe [24], interpreted as due to a cosmological constant or a dark energy component. When combined with the CMB data (which indicate flatness, *i.e.*,  $\Omega_m + \Omega_\Lambda = 1$ ), the best-fit values were  $\Omega_m \approx 0.3$  and  $\Omega_\Lambda \approx 0.7$ . Most results in the literature are consistent with the  $w = -1$  cosmological constant case. A recent study [28] deduced, from a sample of 740 spectroscopically-confirmed SNe Ia, that  $\Omega_m = 0.295 \pm 0.034$  (stat+sym) for an assumed flat  $\Lambda$ CDM model. This is consistent with the latest CMB measurements. Future experiments will aim to set constraints on the cosmic equation of state  $w(z)$ .

### 25.3.3. Cosmic microwave background :

The physics of the CMB is described in detail in the CMB chapter in this volume. Before recombination, the baryons and photons are tightly coupled, and the perturbations oscillate in the potential wells generated primarily by the dark matter perturbations. After decoupling, the baryons are free to collapse into those potential wells. The CMB carries a record of conditions at the time of last scattering, often called primary anisotropies. In addition, it is affected by various processes as it propagates towards us, including the effect of a time-varying gravitational potential (the integrated Sachs–Wolfe effect), gravitational lensing, and scattering from ionized gas at low redshift.

The primary anisotropies, the integrated Sachs–Wolfe effect, and the scattering from a homogeneous distribution of ionized gas, can all be calculated using linear perturbation theory. Available codes include CAMB and CLASS [9], the former widely used embedded within the analysis package CosmoMC [29]. Gravitational lensing is also calculated in these codes. Secondary effects such as inhomogeneities in the reionization process, and scattering from gravitationally-collapsed gas (the Sunyaev–Zeldovich (SZ) effect), require more complicated, and more uncertain, calculations.

The upshot is that the detailed pattern of anisotropies depends on all of the cosmological parameters. In a typical cosmology, the anisotropy power spectrum [usually plotted as  $\ell(\ell+1)C_\ell$ ] features a flat plateau at large angular scales (small  $\ell$ ), followed by a series of oscillatory features at higher angular scales, the first and most prominent being at around one degree ( $\ell \simeq 200$ ). These features, known as acoustic peaks, represent the oscillations of the photon–baryon fluid around the time of decoupling. Some features can be closely related to specific parameters—for instance, the location in multipole space of the set of peaks probes the spatial geometry, while the relative heights of the peaks probe the baryon density—but many other parameters combine to determine the overall shape.

The 2015 data release from the *Planck* satellite [1] gives the most powerful results to date on the spectrum of CMB temperature anisotropies, with a precision determination of the temperature power spectrum to beyond  $\ell = 2000$ . The Atacama Cosmology Telescope (ACT) and South Pole Telescope (SPT) experiments extend these results to higher angular resolution, though without full-sky coverage. *Planck* and the *WMAP* satellite final (9-year) data release [3] give the state of the art in measuring the spectrum of  $E$ -polarization anisotropies and the correlation spectrum between temperature and polarization (those spectra having first been detected by DASI [30]). These are consistent with models based on the parameters we have described, and provide accurate determinations of many of those parameters [2]. Primordial  $B$ -mode polarization has not been

detected (although the gravitational lensing effect on  $B$ -mode has been measured).

The data provide an exquisite measurement of the location of the set of acoustic peaks, determining the angular-diameter distance of the last-scattering surface. In combination with other data this strongly constrains the spatial geometry, in a manner consistent with spatial flatness and excluding significantly-curved Universes. CMB data give a precision measurement of the age of the Universe. The CMB also gives a baryon density consistent with, and at higher precision than, that coming from BBN. It affirms the need for both dark matter and dark energy. It shows no evidence for dynamics of the dark energy, being consistent with a pure cosmological constant ( $w = -1$ ). The density perturbations are consistent with a power-law primordial spectrum, and there is no indication yet of tensor perturbations. The current best-fit for the reionization optical depth from CMB data,  $\tau = 0.066$ , is in line with models of how early structure formation induces reionization.

*Planck* has also made the first all-sky map of the CMB lensing field, which probes the entire matter distribution in the Universe and adds some additional constraining power to the CMB-only data-sets. ACT previously announced the first detection of gravitational lensing of the CMB from the four-point correlation of temperature variations [31]. These measurements agree with the expected effect in the standard cosmology.

#### 25.3.4. Galaxy clustering :

The power spectrum of density perturbations depends on the nature of the dark matter. Within the  $\Lambda$ CDM model, the power spectrum shape depends primarily on the primordial power spectrum and on the combination  $\Omega_m h$ , which determines the horizon scale at matter-radiation equality, with a subdominant dependence on the baryon density. The matter distribution is most easily probed by observing the galaxy distribution, but this must be done with care since the galaxies do not perfectly trace the dark matter distribution. Rather, they are a ‘biased’ tracer of the dark matter. The need to allow for such bias is emphasized by the observation that different types of galaxies show bias with respect to each other. In particular, scale-dependent and stochastic biasing may introduce a systematic effect on the determination of cosmological parameters from redshift surveys. Prior knowledge from simulations of galaxy formation or from gravitational lensing data could help to quantify biasing. Furthermore, the observed 3D galaxy distribution is in redshift space, *i.e.*, the observed redshift is the sum of the Hubble expansion and the line-of-sight peculiar velocity, leading to linear and non-linear dynamical effects that also depend on the cosmological parameters. On the largest length scales, the galaxies are expected to trace the location of the dark matter, except for a constant multiplier  $b$  to the power spectrum, known as the linear bias parameter. On scales smaller than  $20 h^{-1}$  Mpc or so, the clustering pattern is ‘squashed’ in the radial direction due to coherent infall, which depends approximately on the parameter  $\beta \equiv \Omega_m^{0.6}/b$  (on these shorter scales, more complicated forms of biasing are not excluded by the data). On scales of a few  $h^{-1}$  Mpc, there is an effect of elongation along the line of sight (colloquially known as the ‘finger of God’ effect) that depends on the galaxy velocity dispersion.

##### 25.3.4.1. Baryonic acoustic oscillations:

The power spectra of the 2-degree Field (2dF) Galaxy Redshift Survey and the Sloan Digital Sky Survey (SDSS) are well fit by a  $\Lambda$ CDM model and both surveys showed evidence for baryon acoustic oscillations (BAOs) [32,33]. The Baryon Oscillation Spectroscopic Survey (BOSS) of Luminous Red Galaxies (LRGs) in the SDSS found consistency with the dark energy equation of state  $w = -1$  [35]. Similar results for  $w$  were obtained by the WiggleZ survey [36].

##### 25.3.4.2. Redshift distortion:

There is renewed interest in the ‘redshift distortion’ effect. This distortion depends on cosmological parameters [38] via the perturbation growth rate in linear theory  $f(z) = d \ln \delta / d \ln a \approx \Omega^\gamma(z)$ , where  $\gamma \simeq 0.55$  for the  $\Lambda$ CDM model and may be different for modified gravity models. Recent observational results show that by measuring  $f(z)$  it is feasible to constrain  $\gamma$  and rule out certain modified gravity

models [39,40]. We note the degeneracy of the redshift-distortion pattern and the geometric distortion (the so-called Alcock–Paczynski effect [41]), *e.g.*, as illustrated by the WiggleZ survey [42].

##### 25.3.4.3. Limits on neutrino mass from galaxy surveys and other probes:

Large-scale structure data place constraints on  $\Omega_\nu$  due to the neutrino free-streaming effect [47]. Presently there is no clear detection, and upper limits on neutrino mass are commonly estimated by comparing the observed galaxy power spectrum with a four-component model of baryons, cold dark matter, a cosmological constant, and massive neutrinos. Such analyses also assume that the primordial power spectrum is adiabatic, scale-invariant, and Gaussian. Potential systematic effects include biasing of the galaxy distribution and non-linearities of the power spectrum. An upper limit can also be derived from CMB anisotropies alone, while combination with additional cosmological data-sets can improve the results.

Results using a photometric redshift sample of LRGs combined with *WMAP*, BAO, Hubble constant and SNe Ia data gave a 95% confidence upper limit on the total neutrino mass of 0.28 eV [48]. Recent spectroscopic redshift surveys, with more accurate redshifts but fewer galaxies, yield similar upper limits for assumed flat  $\Lambda$ CDM model and additional data-sets: 0.34 eV from BOSS [49] and 0.29 eV from WiggleZ [50].

The *Planck* collaboration [2] derived from only TT+lensing data (see their Table 5), without external data sets, a neutrino mass upper limit of 0.675 eV (95% CL) and  $N_{\text{eff}} = 3.13 \pm 0.31$  (68% CL), in good agreement with the standard value  $N_{\text{eff}} = 3.046$ . When adding external data the upper limit on the neutrino mass is reduced to 0.234 eV, consistent with the above-mentioned pre-Planck results. The Planck result for  $N_{\text{eff}}$  changes little when adding external data.

While the latest cosmological data do not yet constrain the sum of neutrino masses to below 0.2 eV, because the lower limit on neutrino mass from terrestrial experiments is 0.06 eV it appears promising that future cosmological surveys will detect effects from the neutrino mass.

##### 25.3.5. Clustering in the inter-galactic medium :

It is commonly assumed, based on hydrodynamic simulations, that the neutral hydrogen in the inter-galactic medium (IGM) can be related to the underlying mass distribution. It is then possible to estimate the matter power spectrum on scales of a few megaparsecs from the absorption observed in quasar spectra, the so-called Lyman- $\alpha$  forest. The usual procedure is to measure the power spectrum of the transmitted flux, and then to infer the mass power spectrum. Photo-ionization heating by the ultraviolet background radiation and adiabatic cooling by the expansion of the Universe combine to give a simple power-law relation between the gas temperature and the baryon density. It also follows that there is a power-law relation between the optical depth  $\tau$  and  $\rho_b$ . Therefore, the observed flux  $F = \exp(-\tau)$  is strongly correlated with  $\rho_b$ , which itself traces the mass density. The matter and flux power spectra can be related by a biasing function that is calibrated from simulations. The BOSS survey has been used to detect and measure the BAO feature in the Lyman- $\alpha$  forest fluctuation at redshift  $z = 2.4$ , with a result impressively consistent with the standard  $\Lambda$ CDM model [52]. The Lyman- $\alpha$  flux power spectrum has also been used to constrain the nature of dark matter, for example constraining the amount of warm dark matter [53].

##### 25.3.6. Gravitational lensing :

Images of background galaxies are distorted by the gravitational effect of mass variations along the line of sight. Deep gravitational potential wells such as galaxy clusters generate ‘strong lensing’, leading to arcs, arclets and multiple images, while more moderate perturbations give rise to ‘weak lensing’. Weak lensing is now widely used to measure the mass power spectrum in selected regions of the sky (see Ref. 54 for reviews). As the signal is weak, the image of deformed galaxy shapes (the ‘shear map’) must be analyzed statistically to measure the power spectrum, higher moments, and cosmological parameters. There are various systematic effects in the interpretation of weak lensing, *e.g.*, due to atmospheric distortions during observations, the redshift distribution of the background galaxies, the intrinsic correlation of galaxy shapes, and non-linear modeling uncertainties.

The shear measurements are mainly sensitive to a combination of  $\Omega_m$  and the amplitude  $\sigma_8$ . For example, the weak-lensing signal detected by the CFHTLenS Survey (over 154 sq deg in 5 optical bands) yields, for a flat  $\Lambda$ CDM model,  $\sigma_8(\Omega_m/0.30)^{0.5} = 0.71 \pm 0.04$  [55]. The results from the Dark Energy Survey science verification area (over 139 sq deg) give [56]  $\sigma_8(\Omega_m/0.30)^{0.5} = 0.81 \pm 0.06$  (after marginalizing over 3 other cosmological parameters and 7 systematic parameters). See the Dark Energy chapter in this volume for other results for  $\sigma_8$  from other probes.

### 25.3.7. Other probes :

Other probes that have been used to constrain cosmological parameters, but that are not presently competitive in terms of accuracy, are the integrated Sachs–Wolfe effect ([43], [46]) number density or composition of galaxy clusters [51], and galaxy peculiar velocities, which probe the mass fluctuations in the local universe [57].

## 25.4. Bringing observations together

Although it contains two ingredients—dark matter and dark energy—which have not yet been verified by laboratory experiments, the  $\Lambda$ CDM model is almost universally accepted by cosmologists as the best description of the present data. The approximate values of some of the key parameters are  $\Omega_b \approx 0.05$ ,  $\Omega_c \approx 0.25$ ,  $\Omega_\Lambda \approx 0.70$ , and a Hubble constant  $h \approx 0.70$ . The spatial geometry is very close to flat (and usually assumed to be precisely flat), and the initial perturbations Gaussian, adiabatic, and nearly scale-invariant.

The most powerful data source is the CMB, which on its own supports all these main tenets. Values for some parameters, as given in Ref. 2, are reproduced in Table 25.1. These particular results presume a flat Universe. The constraints are somewhat strengthened by adding additional data-sets such as BAO and supernovae, as shown in the Table, though most of the constraining power resides in the CMB data. Similar constraints were previously obtained by the *WMAP* collaboration; the additional precision of *Planck* data versus *WMAP* is only really apparent when considering larger parameter sets.

If the assumption of spatial flatness is lifted, it turns out that the CMB on its own only weakly constrains the spatial curvature, due to a parameter degeneracy in the angular-diameter distance. However, inclusion of other data readily removes this. For example, adding the usual non-CMB data-sets, plus the assumption that the dark energy is a cosmological constant, yields a 68% confidence constraint on  $\Omega_{\text{tot}} \equiv \sum \Omega_i + \Omega_\Lambda = 1.0002 \pm 0.0026$  [2]. Results of this type are normally taken as justifying the restriction to flat cosmologies.

One derived parameter that is very robust is the age of the Universe, since there is a useful coincidence that for a flat Universe the position of the first peak is strongly correlated with the age. The CMB data give  $13.80 \pm 0.04$  Gyr (assuming flatness). This is in good agreement with the ages of the oldest globular clusters and with radioactive dating.

The baryon density  $\Omega_b$  is now measured with high accuracy from CMB data alone, and is consistent with and much more precise than the determination from BBN. The value quoted in the Big Bang Nucleosynthesis chapter in this volume is  $0.021 \leq \Omega_b h^2 \leq 0.024$  (95% confidence).

While  $\Omega_\Lambda$  is measured to be non-zero with very high confidence, there is no evidence of evolution of the dark energy density. As shown in the Dark Energy chapter in this volume, from a compilation of CMB, SN and BAO measurements, assuming a flat universe,  $w = -0.97 \pm 0.05$ , consistent with the cosmological constant case  $w = -1$ . Allowing more complicated forms of dark energy weakens the limits.

The data provide strong support for the main predictions of the simplest inflation models: spatial flatness and adiabatic, Gaussian, nearly scale-invariant density perturbations. But it is disappointing that there is no sign of primordial gravitational waves, with an upper limit  $r < 0.11$  at 95% confidence [58] (weakening if running is allowed). The spectral index is clearly required to be less than one by current data, though the strength of that conclusion can weaken if additional parameters are included in the model fits.

**Table 25.1:** Parameter constraints reproduced from Ref. 2 (Table 4), with some additional rounding. Both columns assume the  $\Lambda$ CDM cosmology with a power-law initial spectrum, no tensors, spatial flatness, a cosmological constant as dark energy, and the sum of neutrino masses fixed to 0.06eV. Above the line are the six parameter combinations actually fit to the data ( $\theta_{\text{MC}}$  is a measure of the sound horizon at last scattering); those below the line are derived from these. The first column uses *Planck* primary CMB data, restricting polarization data to low multipoles as currently recommended by the Planck collaboration, plus the *Planck* measurement of CMB lensing. This column gives our present recommended values. The second column adds additional data and is included to show that the effect of its inclusion is modest; the extra data are the Hubble parameter, BAO measurements from the SDSS, BOSS, and 6dF surveys, and supernova constraints from the JLA analysis. The perturbation amplitude  $\Delta_{\mathcal{R}}^2$  (denoted  $A_s$  in the original paper) is specified at the scale  $0.05 \text{ Mpc}^{-1}$ . Uncertainties are shown at 68% confidence.

	<i>Planck</i> TT+lowP+lensing	<i>Planck</i> TT+lowP+lensing+ext
$\Omega_b h^2$	$0.02226 \pm 0.00023$	$0.02227 \pm 0.00020$
$\Omega_c h^2$	$0.1186 \pm 0.0020$	$0.1184 \pm 0.0012$
$100 \theta_{\text{MC}}$	$1.0410 \pm 0.0005$	$1.0411 \pm 0.0004$
$n_s$	$0.968 \pm 0.006$	$0.968 \pm 0.004$
$\tau$	$0.066 \pm 0.016$	$0.067 \pm 0.013$
$\ln(10^{10} \Delta_{\mathcal{R}}^2)$	$3.062 \pm 0.029$	$3.064 \pm 0.024$
$h$	$0.678 \pm 0.009$	$0.679 \pm 0.006$
$\sigma_8$	$0.815 \pm 0.009$	$0.815 \pm 0.009$
$\Omega_m$	$0.308 \pm 0.012$	$0.306 \pm 0.007$
$\Omega_\Lambda$	$0.692 \pm 0.012$	$0.694 \pm 0.007$

Tests have been made for various types of non-Gaussianity, a particular example being a parameter  $f_{\text{NL}}$  that measures a quadratic contribution to the perturbations. Various non-Gaussian shapes are possible (see Ref. 59 for details), and current constraints on the popular ‘local’, ‘equilateral’, and ‘orthogonal’ types combining temperature and polarization data are  $f_{\text{NL}}^{\text{local}} = 1 \pm 5$ ,  $f_{\text{NL}}^{\text{equil}} = -4 \pm 43$ , and  $f_{\text{NL}}^{\text{ortho}} = -26 \pm 21$  respectively (these look weak, but prominent non-Gaussianity requires the product  $f_{\text{NL}} \Delta_{\mathcal{R}}$  to be large, and  $\Delta_{\mathcal{R}}$  is of order  $10^{-5}$ ). Clearly none of these give any indication of primordial non-gaussianity.

## 25.5. Outlook for the future

The concordance model is now well established, and there seems little room left for any dramatic revision of this paradigm. A measure of the strength of that statement is how difficult it has proven to formulate convincing alternatives.

Should there indeed be no major revision of the current paradigm, we can expect future developments to take one of two directions. Either the existing parameter set will continue to prove sufficient to explain the data, with the parameters subject to ever-tightening constraints, or it will become necessary to deploy new parameters. The latter outcome would be very much the more interesting, offering a route towards understanding new physical processes relevant to the cosmological evolution. There are many possibilities on offer for striking discoveries, for example:

- the cosmological effects of a neutrino mass may be unambiguously detected, shedding light on fundamental neutrino properties;
- detection of primordial non-Gaussianities would indicate that

non-linear processes influence the perturbation generation mechanism;

- detection of variation in the dark-energy density (*i.e.*,  $w \neq -1$ ) would provide much-needed experimental input into its nature.

These provide more than enough motivation for continued efforts to test the cosmological model and improve its accuracy. Over the coming years, there are a wide range of new observations that will bring further precision to cosmological studies. Indeed, there are far too many for us to be able to mention them all here, and so we will just highlight a few areas.

The CMB observations will improve in several directions. A current frontier is the study of polarization, for which power spectrum measurements have now been made by several experiments. Detection of primordial  $B$ -mode anisotropies is the next major goal and a variety of projects are targeting this, though theory gives little guidance as to the likely signal level.

An impressive array of cosmology surveys are already operational, under construction, or proposed, including the ground-based Dark Energy Survey (DES), Hyper Suprime Camera (HSC) and Large Synoptic Survey Telescope (LSST), imaging surveys, spectroscopic surveys such as the Dark Energy Spectroscopic Instrument (DESI), and space missions Euclid and the Wide-Field Infrared Survey (WFIRST).

An exciting area for the future is radio surveys of the redshifted 21-cm line of hydrogen. Because of the intrinsic narrowness of this line, by tuning the bandpass the emission from narrow redshift slices of the Universe will be measured to extremely high redshift, probing the details of the reionization process at redshifts up to perhaps 20, as well as measuring large scale features such as the BAOs. LOFAR is the first instrument able to do this and has begun its operations. In the longer term, the Square Kilometre Array (SKA) will take these studies to a precision level.

The development of the first precision cosmological model is a major achievement. However, it is important not to lose sight of the motivation for developing such a model, which is to understand the underlying physical processes at work governing the Universe's evolution. On that angle, progress has been much less dramatic. For instance, there are many proposals for the nature of the dark matter, but no consensus as to which is correct. The nature of the dark energy remains a mystery. Even the baryon density, now measured to an accuracy of a percent, lacks an underlying theory able to predict it within orders of magnitude. Precision cosmology may have arrived, but at present many key questions remain to motivate and challenge the cosmology community.

## References:

1. Planck Collab. 2015 Results I, *Astron. & Astrophys.*, **arXiv:1502.01582v2**.
2. Planck Collab. 2015 Results XIII, *Astron. & Astrophys.*, **arXiv:1502.01589v2**.
3. C. Bennett *et al.*, *Astrophys. J. Supp.* **208**, 20 (2013).
4. G. Hinshaw *et al.*, *Astrophys. J. Supp.* **208**, 19 (2013).
5. S. Fukuda *et al.*, *Phys. Rev. Lett.* **85**, 3999 (2000); Q.R. Ahmad *et al.*, *Phys. Rev. Lett.* **87**, 071301 (2001).
6. E.W. Kolb and M.S. Turner, *The Early Universe*, Addison-Wesley (Redwood City, 1990).
7. D.H. Lyth and A.R. Liddle, *The Primordial Density Perturbation*, Cambridge University Press (2009).
8. A.R. Liddle and D.H. Lyth, *Phys. Lett.* **B291**, 391 (1992).
9. A. Lewis, A. Challinor, and A. Lasenby, *Astrophys. J.* **538**, 473 (2000); D. Blas, J. Lesgourgues, and T. Tram, *Journal of Cosmology and Astroparticle Physics*, **1107**, 034 (2011).
10. D. Fixsen, *Astrophys. J.* **707**, 916 (2009).
11. M. Hobson *et al.* (eds). *Bayesian Methods in Cosmology*, Cambridge University Press (2009).
12. A. Kosowsky and M.S. Turner, *Phys. Rev.* **D52**, 1739 (1995).
13. K.A. Malik and D. Wands, *Phys. Reports* **475**, 1 (2009).
14. Planck Collab. 2013 Results XXV, *Astron. & Astrophys.* **571**, A25 (2014).
15. D.H. Lyth and D. Wands, *Phys. Lett.* **B524**, 5 (2002); K. Enqvist and M.S. Sloth, *Nucl. Phys.* **B626**, 395 (2002); T. Moroi and T. Takahashi, *Phys. Lett.* **B522**, 215 (2001).
16. G. Mangano *et al.*, *Nucl. Phys.* **B729**, 221 (2005).
17. S. Riemer-Sørensen, D. Parkinson, and T.M. Davis, *Publications of Astronomical Society of the Pacific* **30**, e029 (2013).
18. J.K. Webb *et al.*, *Phys. Rev. Lett.* **107**, 191101 (2011); J.A. King *et al.*, *MNRAS* **422**, 3370 (2012); P. Molaro *et al.*, *Astron. & Astrophys.* **555**, 68 (2013).
19. Planck Collab. 2015 Results XVIII, *Astron. & Astrophys.*, **arXiv:1502.01593v1**.
20. W.L. Freedman *et al.*, *Astrophys. J.* **553**, 47 (2001).
21. A.G. Riess *et al.*, *Astrophys. J.* **730**, 119 (2011).
22. E.M.L. Humphreys *et al.*, *Astrophys. J.*, **775**, 13 (2013).
23. B. Leibundgut, *Ann. Rev. Astron. Astrophys.* **39**, 67 (2001).
24. A.G. Riess *et al.*, *Astron. J.* **116**, 1009 (1998); P. Garnavich *et al.*, *Astrophys. J.* **509**, 74 (1998); S. Perlmutter *et al.*, *Astrophys. J.* **517**, 565 (1999).
25. A. Conley *et al.*, *Astrophys. J. Supp.* **192**, 1 (2011); M. Sullivan *et al.*, *Astrophys. J.* **737**, 102 (2011).
26. M. Kowalski *et al.*, *Astrophys. J.* **686**, 749 (2008).
27. R. Kessler *et al.*, *Astrophys. J. Supp.* **185**, 32 (2009).
28. M. Betoule *et al.*, *Astron. & Astrophys.*, **568**, 22 (2014).
29. A. Lewis and S. Bridle, *Phys. Rev.* **D66**, 103511 (2002).
30. J. Kovac *et al.*, *Nature* **420**, 772 (2002).
31. S. Das *et al.*, *Phys. Rev. Lett.* **107**, 021301 (2011).
32. D. Eisenstein *et al.*, *Astrophys. J.* **633**, 560 (2005).
33. S. Cole *et al.*, *MNRAS* **362**, 505 (2005).
34. A. Sanchez *et al.*, *MNRAS* **433**, 1202 (2013).
35. L. Anderson *et al.*, *MNRAS* **441**, 24 (2014).
36. D. Parkinson *et al.*, *Phys. Rev.* **D86**, 103518 (2012).
37. C. Blake *et al.*, *MNRAS* **418**, 1707 (2011).
38. N. Kaiser, *MNRAS* **227**, 1 (1987).
39. L. Guzzo *et al.*, *Nature* **451**, 541 (2008).
40. A. Nusser and M. Davis, *Astrophys. J.* **736**, 93 (2011).
41. C. Alcock and B. Paczynski, *Nature* **281**, 358 (1979).
42. C. Blake *et al.*, *MNRAS* **425**, 405 (2012).
43. R.G. Crittenden and N. Turok, *Phys. Rev. Lett.* **75**, 2642 (1995).
44. T. Giannantonio *et al.*, *MNRAS* **426**, 258 (2012).
45. Planck Collab. 2015 Results XI, *Astron. & Astrophys.*, **arXiv:1507.02704v1**.
46. Planck Collab. 2013 Results XIX, *Astron. & Astrophys.* **571**, A19 (2014).
47. J. Lesgourgues and S. Pastor, *Phys. Reports* **429**, 307 (2006).
48. S. Thomas, F.B. Abdalla, and O. Lahav, *Phys. Rev. Lett.* **105**, 031301 (2010).
49. G.-B. Zhao *et al.*, *MNRAS* **436**, 2038 (2013).
50. S. Riemer-Sørensen *et al.*, *Phys. Rev.* **D85**, 081101 (2012).
51. Planck Collab. 2013 Results XX, *Astron. & Astrophys.* **571**, A20 (2014).
52. A. Slosar *et al.*, *JCAP* **1404**, 026 (2014).
53. M. Viel *et al.*, *Phys. Rev.* **D88**, 043502 (2013).
54. A. Refregier, *Ann. Rev. Astron. Astrophys.* **41**, 645 (2003); R. Massey *et al.*, *Nature* **445**, 286 (2007); H. Hoekstra and B. Jain, *Ann. Rev. Nucl. and Part. Sci.* **58**, 99 (2008).
55. M. Kilbinger *et al.*, *MNRAS* **430**, 2200 (2013); C. Heymans *et al.*, *MNRAS* **427**, 146 (2012); C. Heymans *et al.*, *MNRAS* **432**, 2433 (2013).
56. Dark Energy Survey Collab., submitted to *MNRAS* (2015), **arXiv:1507.05552**.
57. A. Dekel, *Ann. Rev. Astron. Astrophys.* **32**, 371 (1994).
58. Planck Collab. 2015 Results XX, *Astron. & Astrophys.*, **arXiv:1502.02114**.
59. Planck Collab. 2015 Results XVII, *Astron. & Astrophys.*, **arXiv:1502.01592**.



## 26. DARK MATTER

Revised September 2015 by M. Drees (Bonn University) and G. Gerbier (Queen's University, Canada).

### 26.1. Theory

#### 26.1.1. Evidence for Dark Matter :

The existence of Dark (*i.e.*, non-luminous and non-absorbing) Matter (DM) is by now well established [1,2]. The earliest, and perhaps still most convincing, evidence for DM came from the observation that various luminous objects (stars, gas clouds, globular clusters, or entire galaxies) move faster than one would expect if they only felt the gravitational attraction of other visible objects. An important example is the measurement of galactic rotation curves. The rotational velocity  $v$  of an object on a stable Keplerian orbit with radius  $r$  around a galaxy scales like  $v(r) \propto \sqrt{M(r)/r}$ , where  $M(r)$  is the mass inside the orbit. If  $r$  lies outside the visible part of the galaxy and mass tracks light, one would expect  $v(r) \propto 1/\sqrt{r}$ . Instead, in most galaxies one finds that  $v$  becomes approximately constant out to the largest values of  $r$  where the rotation curve can be measured; in our own galaxy,  $v \simeq 240$  km/s at the location of our solar system, with little change out to the largest observable radius. This implies the existence of a *dark halo*, with mass density  $\rho(r) \propto 1/r^2$ , *i.e.*,  $M(r) \propto r$ ; at some point  $\rho$  will have to fall off faster (in order to keep the total mass of the galaxy finite), but we do not know at what radius this will happen. This leads to a lower bound on the DM mass density,  $\Omega_{\text{DM}} \gtrsim 0.1$ , where  $\Omega_X \equiv \rho_X/\rho_{\text{crit}}$ ,  $\rho_{\text{crit}}$  being the critical mass density (*i.e.*,  $\Omega_{\text{tot}} = 1$  corresponds to a flat Universe).

The observation of clusters of galaxies tends to give somewhat larger values,  $\Omega_{\text{DM}} \simeq 0.2$ . These observations include measurements of the peculiar velocities of galaxies in the cluster, which are a measure of their potential energy if the cluster is virialized; measurements of the *X-ray* temperature of hot gas in the cluster, which again correlates with the gravitational potential felt by the gas; and—most directly—studies of (weak) gravitational lensing of background galaxies on the cluster.

A particularly compelling example involves the bullet cluster (1E0657-558) which recently (on cosmological time scales) passed through another cluster. As a result, the hot gas forming most of the clusters' baryonic mass was shocked and decelerated, whereas the galaxies in the clusters proceeded on ballistic trajectories. Gravitational lensing shows that most of the total mass also moved ballistically, indicating that DM self-interactions are indeed weak [1].

Many cosmologists consider the existence of old galaxies (detected at redshift  $z \sim 10$ ) to be the strongest argument for the existence of DM. Observations of the cosmic microwave background (CMB) show that density perturbations at  $z \simeq 1,300$  were very small,  $\delta\rho/\rho < 10^{-4}$ . Since (sub-horizon sized) density perturbations grow only in the matter-dominated epoch, and matter domination starts earlier in the presence of DM, density perturbations start to grow earlier when DM is present, therefore allowing an earlier formation of the first galaxies [3].

All these arguments rely on Einsteinian, or Newtonian, gravity. One might thus ask whether the necessity to postulate the existence of DM, sometimes perceived to be ad hoc, could be avoided by modifying the theory of gravity. Indeed, the so-called Modified Newtonian Dynamics (MOND) allows to reproduce many observations on galactic scales, in particular galactic rotation curves, without introducing DM [4]. However, MOND is a purely non-relativistic theory. Attempts to embed it into a relativistic field theory require the existence of additional fields (e.g. a vector field or a second metric), and introduce considerably arbitrariness [4]. Moreover, the correct description of large-scale structure formation seems to require some sort of DM even in these theories [5]. In contrast, successful models of particle DM (see below) can be described in the well established language of quantum field theory, and do not need any modification of General Relativity, which has passed a large number of tests with flying colors [6].

The currently most accurate, if somewhat indirect, determination of  $\Omega_{\text{DM}}$  comes from global fits of cosmological parameters to a variety

of observations; see the Section on Cosmological Parameters for details. For example, using measurements of the anisotropy of the cosmic microwave background (CMB) and of the spatial distribution of galaxies, Ref. 7 finds a density of cold, non-baryonic matter

$$\Omega_{\text{nbm}} h^2 = 0.1186 \pm 0.0020, \quad (26.1)$$

where  $h$  is the Hubble constant in units of 100 km/(s-Mpc). Some part of the baryonic matter density [7],

$$\Omega_{\text{b}} h^2 = 0.02226 \pm 0.00023, \quad (26.2)$$

may well contribute to (baryonic) DM, *e.g.*, MACHOs [8] or cold molecular gas clouds [9].

The DM density in the “neighborhood” of our solar system is also of considerable interest. This was first estimated as early as 1922 by J.H. Jeans, who analyzed the motion of nearby stars transverse to the galactic plane [2]. He concluded that in our galactic neighborhood, the average density of DM must be roughly equal to that of luminous matter (stars, gas, dust). Remarkably enough, a recent estimate finds a quite similar result for the smooth component of the local Dark Matter density [10]:

$$\rho_{\text{DM}}^{\text{local}} = (0.39 \pm 0.03) \cdot (1.2 \pm 0.2) \cdot (1 \pm \delta_{\text{triax}}) \frac{\text{GeV}}{\text{cm}^3}. \quad (26.3)$$

The first term on the right-hand side of Eq. (26.3) gives the average Dark Matter density at a point one solar distance from the center of our galaxy. The second factor accounts for the fact that the baryons in the galactic disk, in which the solar system is located, also increase the local DM density [11]. The third factor in Eq. (26.3) corrects for possible deviations from a purely spherical halo; according to [12],  $\delta_{\text{triax}} \leq 0.2$ . Small substructures (minihaloes, streams) are not likely to change the local DM density significantly [1]. Note that the first factor in Eq. (26.3) has been derived by fitting a complete model of our galaxy to a host of data, including the galactic rotation curve. A “purely local” analysis, only using the motion of nearby stars, gives a consistent result, with an error three times as large [13].

#### 26.1.2. Candidates for Dark Matter :

Analyses of structure formation in the Universe indicate that most DM should be “cold” or “cool”, *i.e.*, should have been non-relativistic at the onset of galaxy formation (when there was a galactic mass inside the causal horizon) [1]. This agrees well with the upper bound [7] on the contribution of light neutrinos to Eq. (26.1),

$$\Omega_{\nu} h^2 \leq 0.0062 \quad 95\% \text{ CL}. \quad (26.4)$$

Candidates for non-baryonic DM in Eq. (26.1) must satisfy several conditions: they must be stable on cosmological time scales (otherwise they would have decayed by now), they must interact very weakly with electromagnetic radiation (otherwise they wouldn't qualify as *dark matter*), and they must have the right relic density. Candidates include primordial black holes, axions, sterile neutrinos, and weakly interacting massive particles (WIMPs).

Primordial black holes must have formed before the era of Big-Bang nucleosynthesis, since otherwise they would have been counted in Eq. (26.2) rather than Eq. (26.1). Such an early creation of a large number of black holes is possible only in certain somewhat contrived cosmological models [14].

The existence of axions [15] was first postulated to solve the strong *CP* problem of QCD; they also occur naturally in superstring theories. They are pseudo Nambu-Goldstone bosons associated with the (mostly) spontaneous breaking of a new global “Peccei-Quinn” (PQ)  $U(1)$  symmetry at scale  $f_a$ ; see the Section on Axions in this *Review* for further details. Although very light, axions would constitute cold DM, since they were produced non-thermally. At temperatures well above the QCD phase transition, the axion is massless, and the axion field can take any value, parameterized by the “misalignment angle”  $\theta_i$ . At  $T \lesssim 1$  GeV, the axion develops a mass  $m_a \sim f_{\pi} m_{\pi} / f_a$  due to instanton effects. Unless the axion field happens to find itself at the minimum of its potential ( $\theta_i = 0$ ), it will begin to oscillate once

$m_a$  becomes comparable to the Hubble parameter  $H$ . These coherent oscillations transform the energy originally stored in the axion field into physical axion quanta. The contribution of this mechanism to the present axion relic density is [1]

$$\Omega_a h^2 = \kappa_a \left( f_a / 10^{12} \text{ GeV} \right)^{1.175} \theta_i^2, \quad (26.5)$$

where the numerical factor  $\kappa_a$  lies roughly between 0.5 and a few. If  $\theta_i \sim \mathcal{O}(1)$ , Eq. (26.5) will saturate Eq. (26.1) for  $f_a \sim 10^{11}$  GeV, comfortably above laboratory and astrophysical constraints [15]; this would correspond to an axion mass around 0.1 meV. However, if the post-inflationary reheat temperature  $T_R > f_a$ , cosmic strings will form during the PQ phase transition at  $T \simeq f_a$ . Their decay will give an additional contribution to  $\Omega_a$ , which is often bigger than that in Eq. (26.5) [1], leading to a smaller preferred value of  $f_a$ , *i.e.*, larger  $m_a$ . On the other hand, values of  $f_a$  near the Planck scale become possible if  $\theta_i$  is for some reason very small.

“Sterile”  $SU(2) \times U(1)_Y$  singlet neutrinos with keV masses [16] could alleviate the “cusp/core problem” [1] of cold DM models. If they were produced non-thermally through mixing with standard neutrinos, they would eventually decay into a standard neutrino and a photon or into three neutrinos.

Weakly interacting massive particles (WIMPs)  $\chi$  are particles with mass roughly between 10 GeV and a few TeV, and with cross sections of approximately weak strength. Within standard cosmology, their present relic density can be calculated reliably if the WIMPs were in thermal and chemical equilibrium with the hot “soup” of Standard Model (SM) particles after inflation. In this case, their density would become exponentially (Boltzmann) suppressed at  $T < m_\chi$ . The WIMPs therefore drop out of thermal equilibrium (“freeze out”) once the rate of reactions that change SM particles into WIMPs or vice versa, which is proportional to the product of the WIMP number density and the WIMP pair annihilation cross section into SM particles  $\sigma_A$  times velocity, becomes smaller than the Hubble expansion rate of the Universe. After freeze out, the co-moving WIMP density remains essentially constant; if the Universe evolved adiabatically after WIMP decoupling, this implies a constant WIMP number to entropy density ratio. Their present relic density is then approximately given by (ignoring logarithmic corrections) [3]

$$\Omega_\chi h^2 \simeq \text{const.} \cdot \frac{T_0^3}{M_{\text{Pl}}^3 \langle \sigma_A v \rangle} \simeq \frac{0.1 \text{ pb} \cdot c}{\langle \sigma_A v \rangle}. \quad (26.6)$$

Here  $T_0$  is the current CMB temperature,  $M_{\text{Pl}}$  is the Planck mass,  $c$  is the speed of light,  $\sigma_A$  is the total annihilation cross section of a pair of WIMPs into SM particles,  $v$  is the relative velocity between the two WIMPs in their cms system, and  $\langle \dots \rangle$  denotes thermal averaging. Freeze out happens at temperature  $T_F \simeq m_\chi/20$  almost independently of the properties of the WIMP. This means that WIMPs are already non-relativistic when they decouple from the thermal plasma; it also implies that Eq. (26.6) is applicable if  $T_R > T_F$ . Notice that the 0.1 pb in Eq. (26.6) contains factors of  $T_0$  and  $M_{\text{Pl}}$ ; it is, therefore, quite intriguing that it “happens” to come out near the typical size of weak interaction cross sections.

The seemingly most obvious WIMP candidate is a heavy neutrino. However, an  $SU(2)$  doublet neutrino will have too small a relic density if its mass exceeds  $M_Z/2$ , as required by LEP data. One can suppress the annihilation cross section, and hence increase the relic density, by postulating mixing between a heavy  $SU(2)$  doublet and some sterile neutrino. However, one also has to require the neutrino to be stable; it is not obvious why a massive neutrino should not be allowed to decay.

The currently best motivated WIMP candidate is, therefore, the lightest superparticle (LSP) in supersymmetric models [17] with exact R-parity (which guarantees the stability of the LSP). Searches for exotic isotopes [18] imply that a stable LSP has to be neutral. This leaves basically two candidates among the superpartners of ordinary particles, a sneutrino, and a neutralino. The negative outcome of various WIMP searches (see below) rules out “ordinary” sneutrinos as primary component of the DM halo of our galaxy. The most widely studied WIMP is therefore the lightest neutralino. Detailed

calculations [1] show that the lightest neutralino will have the desired thermal relic density Eq. (26.1) in at least four distinct regions of parameter space.  $\chi$  could be (mostly) a bino or photino (the superpartner of the  $U(1)_Y$  gauge boson and photon, respectively), if both  $\chi$  and some sleptons have mass below  $\sim 150$  GeV, or if  $m_\chi$  is close to the mass of some sfermion (so that its relic density is reduced through co-annihilation with this sfermion), or if  $2m_\chi$  is close to the mass of the  $CP$ -odd Higgs boson present in supersymmetric models. Finally, Eq. (26.1) can also be satisfied if  $\chi$  has a large higgsino or wino component.

Many non-supersymmetric extensions of the Standard Model also contain viable WIMP candidates [1]. Examples are the lightest  $T$ -odd particle in “Little Higgs” models with conserved  $T$ -parity, or “techni-baryons” in scenarios with an additional, strongly interacting (“technicolor” or similar) gauge group.

Although thermally produced WIMPs are attractive DM candidates because their relic density naturally has at least the right order of magnitude, non-thermal production mechanisms have also been suggested, *e.g.*, LSP production from the decay of some moduli fields [19], from the decay of the inflaton [20], or from the decay of “ $Q$ -balls” (non-topological solitons) formed in the wake of Affleck-Dine baryogenesis [21]. Although LSPs from these sources are typically highly relativistic when produced, they quickly achieve kinetic (but not chemical) equilibrium if  $T_R$  exceeds a few MeV [22] (but stays below  $m_\chi/20$ ). They therefore also contribute to cold DM. Finally, if the WIMPs aren’t their own antiparticles, an asymmetry between WIMPs and antiWIMPs might have been created in the early Universe, possibly by the same (unknown) mechanism that created the baryon antibaryon asymmetry. In such “asymmetric DM” models [23] the WIMP antiWIMP annihilation cross section  $\langle \sigma_{A\bar{v}} \rangle$  should be significantly larger than  $0.1 \text{ pb} \cdot c$ , cf Eq. (26.6).

The absence of signals at the LHC for physics beyond the Standard Model, as well as the discovery of an SM-like Higgs boson with mass near 125 GeV, constrains many well-motivated WIMP models. For example, in constrained versions of the minimal supersymmetrized Standard Model (MSSM) both the absence of supersymmetric signals and the relatively large mass of the Higgs boson favor larger WIMP masses and lower scattering cross sections on nucleons. However, constraints from “new physics” searches apply most directly to strongly interacting particles. Many WIMP models therefore can still accommodate a viable WIMP for a wide range of masses. For example, in supersymmetric models where the bino mass is not related to the other gaugino masses a bino with mass as small as 15 GeV can still have the correct thermal relic density [24]. Even lighter supersymmetric WIMPs can be realized in models with extended Higgs sector [25].

The lack of signals at the LHC may have weakened the argument for WIMPs being embedded in a larger theory that addresses the hierarchy problem. This, and the increasingly stronger limits from direct and indirect WIMP searches (see below), has spawned a plethora of new models of particle DM. For example, particles with masses in the MeV to GeV range still naturally form cold DM, but are difficult to detect with current methods. These models typically require rather light “mediator” particles in order to achieve the correct thermal relic density. Light bosons coupling to (possibly quite heavy) DM particles have also been invoked in order to greatly increase the annihilation cross section of the latter at small velocities, through the so-called Sommerfeld enhancement [26]. Several collider and fixed target experiments have searched for such light mediators, but no signal has been found [27].

Another mechanism to achieve the correct thermal relic density is based on  $2 \leftrightarrow 3$  reactions purely within the dark sector. This requires quite large self interactions between the DM particles, which have therefore been dubbed SIMPs (strongly interacting massive particles) [28]. The SIMP-SIMP elastic scattering cross section  $\sigma$  might even be large enough to affect cosmological structure formation, which roughly requires  $\sigma/m_\chi > 0.1 \text{ b/GeV}$ , where  $m_\chi$  is the mass of the SIMP; this is considerably larger than the elastic scattering cross section of protons. Scalar SIMPs could interact with ordinary matter via Higgs exchange.

Primary black holes (as MACHOs), axions, sterile neutrinos, and WIMPs are all (in principle) detectable with present or near-future technology (see below). There are also particle physics DM candidates which currently seem almost impossible to detect, unless they decay; the present lower limit on their lifetime is of order  $10^{25}$  to  $10^{26}$  s for 100 GeV particles. These include the gravitino (the spin-3/2 superpartner of the graviton), states from the “hidden sector” thought responsible for supersymmetry breaking, and the axino (the spin-1/2 superpartner of the axion) [1].

## 26.2. Experimental detection of Dark Matter

### 26.2.1. The case of baryonic matter in our galaxy :

The search for hidden galactic baryonic matter in the form of MACHOs has been initiated following the suggestion that they may represent a large part of the galactic DM and could be detected through the microlensing effect [8]. The MACHO, EROS, and OGLE collaborations have performed a program of observation of such objects by monitoring the luminosity of millions of stars in the Large and Small Magellanic Clouds for several years. EROS concluded that MACHOs cannot contribute more than 8% to the mass of the galactic halo [29], while MACHO observed a signal at 0.4 solar mass and put an upper limit of 40%. Overall, this strengthens the need for non-baryonic DM, also supported by the arguments developed above.

### 26.2.2. Axion searches :

Axions can be detected by looking for  $a \rightarrow \gamma$  conversion in a strong magnetic field [1]. Such a conversion proceeds through the loop-induced  $a\gamma\gamma$  coupling, whose strength  $g_{a\gamma\gamma}$  is an important parameter of axion models. There is currently only one experiment searching for axionic DM: the ADMX experiment [30], originally situated at the LLNL in California but now running at the University of Washington, started taking data in the first half of 1996. It employs a high quality cavity, whose “Q factor” enhances the conversion rate on resonance, *i.e.*, for  $m_a(c^2 + v_a^2/2) = \hbar\omega_{\text{res}}$ . One then needs to scan the resonance frequency in order to cover a significant range in  $m_a$  or, equivalently,  $f_a$ . ADMX now uses SQUIDS as first-stage amplifiers; their extremely low noise temperature (1.2 K) enhances the conversion signal. Published results [31], combining data taken with conventional amplifiers and SQUIDS, exclude axions with mass between 1.9 and  $3.53 \mu\text{eV}$ , corresponding to  $f_a \simeq 4 \cdot 10^{13}$  GeV, for an assumed local DM density of  $0.45 \text{ GeV}/\text{cm}^3$ , if  $g_{a\gamma\gamma}$  is near the upper end of the theoretically expected range. About five times better limits on  $g_{a\gamma\gamma}$  were achieved [32] for  $1.98 \mu\text{eV} \leq m_a \leq 2.18 \mu\text{eV}$  as well as for  $3.3 \mu\text{eV} \leq m_a \leq 3.65 \mu\text{eV}$ , if a large fraction of the local DM density is due to a single flow of axions with very low velocity dispersion. The ADMX experiment is being upgraded by reducing the cavity and SQUID temperature from the current 1.2 K to about 0.1 K. This should increase the frequency scanning speed for given sensitivity by more than two orders of magnitude, or increase the sensitivity for fixed observation time.

### 26.2.3. Searches for keV Neutrinos :

Relic keV neutrinos  $\nu_s$  can only be detected if they mix with the ordinary neutrinos. This mixing leads to radiative  $\nu_s \rightarrow \nu\gamma$  decays, with lifetime  $\tau_{\nu_s} \simeq 1.8 \cdot 10^{21} \text{ s} \cdot (\sin\theta)^{-2} \cdot (1 \text{ keV}/m_{\nu_s})^5$ , where  $\theta$  is the mixing angle [16]. This gives rise to a flux of mono-energetic photons with  $E_\gamma = m_{\nu_s}/2$ , which might be observable by *X-ray* satellites. In the simplest case the relic  $\nu_s$  are produced only by oscillations of standard neutrinos. Assuming that all lepton-antilepton asymmetries are well below  $10^{-3}$ , the  $\nu_s$  relic density can then be computed uniquely in terms of the mixing angle  $\theta$  and the mass  $m_{\nu_s}$ . The combination of lower bounds on  $m_{\nu_s}$  from analyses of structure formation (in particular, the Ly $\alpha$  “forest”) and upper bounds on *X-ray* fluxes from various (clusters of) galaxies exclude this scenario if  $\nu_s$  forms all of DM. This conclusion can be evaded if  $\nu_s$  forms only part of DM, and/or if there is a lepton asymmetry  $\geq 10^{-3}$  (*i.e.* some 7 orders of magnitude above the observed baryon-antibaryon asymmetry), and/or if there is an additional source of  $\nu_s$  production in the early Universe, *e.g.* from the decay of heavier particles [16].

Recently some evidence for a weak *X-ray* line at  $\sim 3.5 \text{ keV}$  has been found in data released by the XMM-Newton satellite. Although this has been interpreted in terms of decaying keV DM particles, *e.g.* sterile neutrinos with mass  $m_{\nu_s} \simeq 7 \text{ keV}$ , it might also be due to certain inner-shell transitions of highly ionized K atoms [33].

### 26.2.4. Basics of direct WIMP search :

As stated above, WIMPs should be gravitationally trapped inside galaxies and should have the adequate density profile to account for the observed rotational curves. These two constraints determine the main features of experimental detection of WIMPs, which have been detailed in the reviews in [1].

Their mean velocity inside our galaxy relative to its center is expected to be similar to that of stars, *i.e.*, a few hundred kilometers per second at the location of our solar system. For these velocities, WIMPs interact with ordinary matter through elastic scattering on nuclei. With expected WIMP masses in the range 10 GeV to 10 TeV, typical nuclear recoil energies are of order of 1 to 100 keV.

The shape of the nuclear recoil spectrum results from a convolution of the WIMP velocity distribution, usually taken as a Maxwellian distribution in the galactic rest frame, shifted into the Earth rest frame, with the angular scattering distribution, which is isotropic to first approximation but forward-peaked for high nuclear mass (typically higher than Ge mass) due to the nuclear form factor. Overall, this results in a roughly exponential spectrum. The higher the WIMP mass, the higher the mean value of the exponential. This points to the need for low nuclear recoil energy threshold detectors.

On the other hand, expected interaction rates depend on the product of the local WIMP flux and the interaction cross section. The first term is fixed by the local density of dark matter, taken as  $0.39 \text{ GeV}/\text{cm}^3$  [see Eq. (26.3)], the mean WIMP velocity, typically 220 km/s, the galactic escape velocity, typically 544 km/s [34] and the mass of the WIMP. The expected interaction rate then mainly depends on two unknowns, the mass and cross section of the WIMP (with some uncertainty [10] due to the halo model). This is why the experimental observable, which is basically the scattering rate as a function of energy, is usually expressed as a contour in the WIMP mass-cross section plane.

The cross section depends on the nature of the couplings. For non-relativistic WIMPs, one in general has to distinguish spin-independent and spin-dependent couplings. The former can involve scalar and vector WIMP and nucleon currents (vector currents are absent for Majorana WIMPs, *e.g.*, the neutralino), while the latter involve axial vector currents (and obviously only exist if  $\chi$  carries spin). Due to coherence effects, the spin-independent cross section scales approximately as the square of the mass of the nucleus, so higher mass nuclei, from Ge to Xe, are preferred for this search. For spin-dependent coupling, the cross section depends on the nuclear spin factor; used target nuclei include  $^{19}\text{F}$ ,  $^{23}\text{Na}$ ,  $^{73}\text{Ge}$ ,  $^{127}\text{I}$ ,  $^{129}\text{Xe}$ ,  $^{131}\text{Xe}$ , and  $^{133}\text{Cs}$ .

Cross sections calculated in MSSM models [35] induce rates of at most  $1 \text{ evt day}^{-1} \text{ kg}^{-1}$  of detector, much lower than the usual radioactive backgrounds. This indicates the need for underground laboratories to protect against cosmic ray induced backgrounds, and for the selection of extremely radio-pure materials.

The typical shape of exclusion contours can be anticipated from this discussion: at low WIMP mass, the sensitivity drops because of the detector energy threshold, whereas at high masses, the sensitivity also decreases because, for a fixed mass density, the WIMP flux decreases  $\propto 1/m_\chi$ . The sensitivity is best for WIMP masses near the mass of the recoiling nucleus.

Two important points are to be kept in mind when comparing exclusion curves from various experiments between them or with positive indications of a signal.

For an experiment with a fixed nuclear recoil energy threshold, the lower is the considered WIMP mass, the lower is the fraction of the spectrum to which the experiment is sensitive. This fraction may be extremely small in some cases. For instance CoGeNT [36], using a Germanium detector with an energy threshold of around 2 keV, is sensitive to about 10 % of the total recoil spectrum of a 7 GeV

WIMP, while for XENON100 [37], using a liquid Xenon detector with a threshold of 8.4 keV, this fraction is only 0.05 % (that is the extreme tail of the distribution), for the same WIMP mass. The two experiments are then sensitive to very different parts of the WIMP velocity distribution.

A second important point to consider is the energy resolution of the detector. Again at low WIMP mass, the expected roughly exponential spectrum is very steep and when the characteristic energy of the exponential becomes of the same order as the energy resolution, the energy smearing becomes important. In particular, a significant fraction of the expected spectrum below effective threshold is smeared above threshold, increasing artificially the sensitivity. For instance, a Xenon detector with a threshold of 8 keV and infinitely good resolution is actually insensitive to a 7 GeV mass WIMP, because the expected energy distribution has a cut-off at roughly 5 keV. When folding in the experimental resolution of XENON100 (corresponding to a photostatistics of 0.5 photoelectron per keV), then around 1 % of the signal is smeared above 5 keV and 0.05 % above 8 keV. Setting reliable cross section limits in this mass range thus requires a complete understanding of the response of the detector at energies well below the nominal threshold.

In order to homogenize the reliability of the presented exclusion curves, and save the reader the trouble of performing tedious calculations, we propose to set cross section limits only for WIMP mass above a “WIMP safe” minimal mass value defined as the maximum of 1) the mass where the increase of sensitivity from infinite resolution to actual experimental resolution is not more than a factor two, and 2) the mass where the experiment is sensitive to at least 1 % of the total WIMP signal recoil spectrum. These recommendations are irrespective of the content of the experimental data obtained by the experiments.

Two experimental signatures are predicted for WIMP signals. One is a strong daily forward/backward asymmetry of the nuclear recoil direction, due to the alternate sweeping of the WIMP cloud by the rotating Earth. Detection of this effect requires gaseous detectors or anisotropic response scintillators (stilbene). The second is a few percent annual modulation of the recoil rate due to the Earth speed adding to or subtracting from the speed of the Sun. This tiny effect can only be detected with large masses; nuclear recoil identification should also be performed, as the otherwise much larger background may also be subject to seasonal modulation.

#### 26.2.5. Status and prospects of direct WIMP searches :

Given the intense activity of the field, readers interested in more details than the ones given below may refer to [1], to presentations at recent conferences [30] and to the previous versions of this review.

The first searches have been performed with ultra-pure semiconductors installed in pure lead and copper shields in underground environments. Combining a priori excellent energy resolutions and very pure detector material, they produced the first limits on WIMP searches (Heidelberg-Moscow, IGEX, COSME-II, HDMS) [1]. Planned experiments using several tens of kg to a ton of Germanium run at liquid nitrogen temperature (designed for double-beta decay search) – GERDA, MAJORANA – are based in addition on passive reduction of the external and internal electromagnetic and neutron background by using Point Contact detectors (discussed below), minimal detector housing, close electronics, pulse shape discrimination and large liquid nitrogen or argon shields. Their sensitivity to WIMP interactions will depend on their ability to lower the energy threshold sufficiently, while keeping the background rate small.

Development of so called Point Contact Germanium detectors, with a very small capacitance allowed one to reach sub-keV thresholds, though performance seems to stall now at around 400 eV. The CoGeNT collaboration was first operating a single 440 g Germanium detector with an effective threshold of 400 eV in the Soudan Underground Laboratory for 56 days [36]. The originally quite large “not understood” excess at low energy claimed a couple of years ago has been reduced by a more careful treatment of surface events. Given the still substantial systematic uncertainties attached to this background, the remaining excess, if any, is not significant.

The annual modulation compatible with a dark matter signal claimed by CoGeNT also fell short. Two unpublished papers [38] on data taken over a period of 3.4 years were written by different CoGeNT authors. The more frequently cited one states a modest 2.2 sigma excess, with the comment “However, its phase is compatible with that predicted by halo simulations, and observed by DAMA/LIBRA”, thus still not giving up a Dark Matter interpretation. The other concludes that “the Null (no-WIMP) hypothesis is only excluded at less than 2 sigma”, so does not claim any signal. Finally, an independent analysis by J. Davis *et al.* [39] concludes that “the CoGeNT data show a preference for light dark matter recoils at less than 1 sigma”.

The CDEX collaboration has operated also a single Point Contact detector in the Jinping underground laboratory, with a 475 eV threshold and a background rate too high to lead to a competitive limit. Their next step is CDEX-10, an array with a total mass of 10 kg, planned to be immersed in a ton-scale liquid argon chamber as active shield.

In order to make progress in the reliability of any claimed signal, active background rejection and signal identification questions have to be addressed. Active background rejection in detectors relies on the relatively small ionization in nuclear recoils due to their low velocity. This induces a reduction (“quenching”) of the ionization/scintillation signal for nuclear recoil signal events relative to  $e$  or  $\gamma$  induced backgrounds of the same energy. Energies calibrated with gamma sources are then called “electron equivalent energies” (keVee unit, used below). This effect has been both calculated and measured [1]. It is exploited in cryogenic detectors described later. In scintillation detectors, it induces in addition a difference in decay times of pulses induced by  $e/\gamma$  events vs nuclear recoils. In most cases, due to the limited resolution and discrimination power of this technique at low energies, this effect allows only a statistical background rejection. It has been used in NaI(Tl) (DAMA, LIBRA, NAIAD, Saclay NaI), in CsI(Tl) (KIMS), and Xe (ZEPLIN-I) [1,30]. In liquid argon, pulse shape discrimination applied to the pulse of primary scintillation light is particularly efficient and allows an event by event discrimination, however, at some high energy, roughly above 40 keVee (see the DarkSide50 result later in this review).

The DAMA collaboration has reported results from a total of 7 years exposure with the LIBRA phase involving 250 kg of detectors, plus the earlier 6 years exposure of the original DAMA/NaI experiment with 100 kg of detectors [40], for a cumulated exposure of 1.33 t.y. They observe an annual modulation of the signal in the 2 to 6 keVee bin, with the expected period (1 year) and phase (maximum around June 2), at 9.3  $\sigma$  level. If interpreted within the standard halo model described above, two possible solutions have been proposed: a WIMP with  $m_\chi \simeq 50$  GeV and  $\sigma_{\chi p} \simeq 7 \cdot 10^{-6}$  pb (central values) or at low mass, in the 6 to 10 GeV range with  $\sigma_{\chi p} \sim 10^{-3}$  pb; the cross section could be somewhat lower if there is a significant channeling effect [1].

Interpreting these observations as positive WIMP signal raises several issues of internal consistency. First, the proposed WIMP solutions would induce a sizeable fraction of nuclear recoils in the total measured rate in the 2 to 6 keVee bin. No pulse shape analysis has been reported by the authors to check whether the unmodulated signal was detectable this way. Secondly, the residual  $e/\gamma$ -induced background, inferred by subtracting the signal predicted by the WIMP interpretation from the data, has an unexpected shape [41], starting near zero at threshold and quickly rising to reach its maximum near 3 to 3.5 keVee; from general arguments one would expect the background (e.g. due to electronic noise) to increase towards the threshold. Finally, the amplitude of the annual modulation shows a somewhat troublesome tendency to decrease with time. The original DAMA data, taken 1995 to 2001, gave an amplitude of the modulation of  $20.0 \pm 3.2$  in units of  $10^{-3}$  counts/(kg.day.keVee), in the 2-6 keVee bin. During the first phase of DAMA/LIBRA, covering data taken between 2003 and 2007, this amplitude became  $10.7 \pm 1.9$ , and in the second phase of DAMA/LIBRA, covering data taken between 2007 and 2009, it further decreased to  $8.5 \pm 2.2$ . The ratio of amplitudes inferred from the DAMA/LIBRA phase 2 and original DAMA data is  $0.43 \pm 0.13$ , differing from the expected value of 1 by more than 4 standard deviations. The results for the DAMA/LIBRA phase 2 have been calculated by us using published results for the earlier data

alone [42] as well as for the grand total [43]. Similar conclusions can be drawn from analyses of the 2-4 and 2-5 keVee bins.

The two last years have seen a growing number of projects using NaI(Tl) scintillator (ANAIS, KIMS, SABRE, DM-ICE), some of them likely to reach the maturity to test the DAMA/LIBRA claim. Two of them (ANAIS, KIMS) profit of new crystals delivered by a supplier independent of DAMA, showing a light yield around two times higher than the ones previously used. This opens the possibility of a significant nuclear recoil-electron recoil discrimination at energies down to 2 keV. The KIMS team states that, under the hypothesis that the DAMA signal originates from nuclear recoils, an equivalent sensitivity to the DAMA/LIBRA solutions could be tested by operating 100 kg of detectors during one year [44].

DM-ICE runs detectors within the IceCube neutrino telescope. They found an unexpected phosphorescence with a decay time of about five seconds after the primary scintillation light induced by cosmic ray muons [45]. Their conclusion is this effect of long tail of single photoelectrons cannot mimic the DAMA effect.

SABRE also contemplates the possibility to run detectors in the Southern hemisphere (in ANDES, a project for an underground laboratory in a road tunnel connecting Argentina and Chile, or STAWELL in Australia, in a gold mine 240 km west of Melbourne), in order to test for a possible shift of the phase of the annual modulation. Such a shift would be expected if the modulation is somehow related to the seasons on Earth, whereas a WIMP induced annual modulation should have the same phase in both hemispheres.

As is shown below in the figure summarizing all recent results, under standard assumptions by now many experiments exclude both the high and low mass DAMA/LIBRA solutions. In case of spin independent interactions, these assumptions include a common cross section for WIMPs scattering on protons and neutrons, allowing the direct comparison of results from experiments using detectors with different neutron to proton ratios in their target nuclei. Given that both Germanium and Xenon detectors now exclude the DAMA/LIBRA signal by a large margin under standard assumptions, even allowing independent matrix elements for WIMP interactions with protons and neutrons (“isospin violating” WIMPs) cannot reconcile all experimental results [46]. The large WIMP mass interpretation of the DAMA/LIBRA signal is excluded most directly by results from the KIMS experiment. They operated 12 crystals of CsI(Tl) with a total mass of 104.4 kg in the Yang Yang (renamed CUNP) laboratory in Korea, and gave an upper limit on nuclear recoils present in a 24 t-d exposure [47]. This translates into an upper limit on the cross section roughly two orders of magnitude below that required to explain the DAMA signal by a 60 GeV WIMP, induced by the Iodine nucleus. Here, the direct comparison of experiments is possible as they use the same nucleus. A more recent analysis which extends to low threshold [48] under standard assumptions also excludes most of the low mass solution. On the other hand, no convincing non-WIMP explanation of the annual modulation of the DAMA/LIBRA signal has yet been put forward. For example, it is well known that the cosmic ray muon flux varies with the season, yet this source of background is much too small to explain the effect [49]. After 14 years, the DAMA/LIBRA result therefore continues to inspire (increasingly baroque) theoretical speculations [50].

At mK temperature, the simultaneous measurement of the phonon and ionization signals in semiconductor detectors permits event by event discrimination between nuclear and electronic recoils down to few keV recoil energy. This feature is being used by the CDMS [30] and EDELWEISS [30] collaborations. Surface interactions, exhibiting incomplete charge collection, are an important residual background. Both experiments now use an interleaved ionization read out electrodes scheme in order to control this background.

The total CDMS exposure of 612 kg-d (around 300 kg-d fiducial) from 2011 data using 19 Germanium cryogenic detectors at the Soudan mine has recently been reanalysed with an improved charge-pulse fitting algorithm [51]. It provides a new spin-independent WIMP nucleon cross sections limit at  $1.8 \times 10^{-8}$  pb, at 90% CL for a 60 GeV WIMP and  $1.8 \times 10^{-5}$  pb for a 8.6 GeV WIMP, a factor of two improvement over their previous analysis (see the 2013 version of this

review).

Since March 2012, CDMS has operated 15 IZIP (using interleaved electrode scheme) detectors. A subset of these data, obtained from the detectors giving the lowest threshold, has been analyzed in view of improving the reach at low WIMP mass. A blind multivariate analysis [52] on 577 kg-d found 11 events, out of which 3 events originated from an unanticipatedly malfunctioning detector. With 6 events expected from background, no hint of a WIMP signal is claimed. Improved limits were set in the 4 to 15 GeV region, excluding scenarios favored by positive claims from CoGENT, CRESST, as well as the CMDS Silicon result.

Two results were reported by CDMS from a single detector running in a particular mode allowing an equivalent electron energy threshold of 170 eV for the first result and then recently of 50-70eV [53]. This is obtained by applying a high voltage (70 V) across the electrodes measuring the ionization. The phonons generated by the ionization electrons traveling inside the crystal – the so-called Neganov Luke effect – give a stronger signal than the normal phonon pulse induced by the initial interaction. This amplifies the ionization pulse, but no discrimination between electron and nuclear recoils is possible in this mode. The sensitivity is then fixed by the counting rate at threshold. A significant improvement is obtained at around 3 GeV, around 1 to 2 order of magnitude in sensitivity relative to previous best results. The limit obtained by this detector is better than that of [52] described above for WIMP masses below 6 GeV.

The EDELWEISS collaboration [30] now operates 30 kg of cryogenic Germanium detectors (so-called FID800 detectors, featuring a complete coverage of the crystal with annular electrodes, and better rejection of non-recoil events) in the Laboratoire Souterrain de Modane. Using a subset of currently acquired data from a single detector with an especially low threshold, corresponding to the rather modest exposure of 37 kg-d, a limit close to the one from CDMS with larger exposure was obtained for WIMP masses around 20 GeV [54]. More data are to come and EDELWEISS also plans to operate HV assisted Neganov Luke detectors.

The combined analysis of CDMS and EDELWEISS data [55] still gives the best limit for cryogenic detectors on the SI cross sections for WIMPs masses above 80 GeV.

The cryogenic experiment CRESST [30] in the Gran Sasso laboratory uses the scintillation of  $\text{CaWO}_4$  crystals as second variable for background discrimination. Like many other experiments in the last few years, CRESST puts focus on lowering the energy threshold to access low mass WIMPs. They use a new generation of detectors with improved vetoing of low energy surface events induced by external alpha particles. Results from two single detectors showing good energy resolution, have been published recently [56]. The last one shows an impressively low threshold of 0.3 keVNR, allowing one to set a limit on WIMP-proton cross section for spin independent couplings of  $10^{-2}$  pb at a WIMP mass of 1 GeV, thanks to the presence of Oxygen nuclei in the target. Interestingly, the obtained limit excludes the signals reported by the same collaboration two years before, which are now believed to have been caused by an inadequate description of the background from external alpha particles.

The next stages of solid state detectors are SuperCDMS and EURECA (a combination of EDELWEISS and CRESST). The SuperCDMS project at SNOLAB has been approved by the US funding agencies with the aim of addressing low mass WIMPs. Given that the current limits on cross sections below WIMP mass of 10 GeV are rather high, the mass required to get significant improvements does not need to be large. A target mass of 50 to 200 kg is even sufficient to reach sensitivity somewhat below  $10^{-8}$  pb for the WIMP-proton cross section after about 5 years of running. In order to achieve this goal lower thresholds and the rejection or very low radioactive backgrounds are mandatory. Improving the sensitivity below this level is very difficult for light WIMPs, due to the irreducible background from the elastic scattering of (mostly solar) neutrinos off the target nuclei (the “neutrino floor”) [57]. The SuperCDMS and EURECA collaborations are working on the common operation of all types of detectors in the same cryogenic set-up at SNOLAB, foreseen to be ready to operate around 2019. Calculated sensitivities down to a

WIMP mass of 1 GeV [58] rely on the extrapolation of knowledge of the radioactive background down to 10 eVee and of the quenching factor down to 50 eVNR.

Noble gas detectors are being actively developed. Due to their relatively easy scalability they are likely the first to be able to perform high sensitivity searches for “high mass” WIMPs (with masses above  $\sim 15$  GeV). Dual (liquid and gas) phase detectors allow to measure both the primary scintillation S1 and the ionization electrons drifted through the liquid, amplified in the gas and giving rise to a second scintillation pulse S2. S1 and S2 are used to perform discrimination and 3D position reconstruction within the detector. In the single phase mode (DEAP, XMASS), only S1 is measured and the discrimination is ensured by the pulse shape analysis in the case of Argon and by the self shielding in the case of Xenon.

The suite of XENON-n detectors [30] are operated at the Gran Sasso laboratory. After XENON10, XENON 100 set the first significant sign of supremacy of liquid noble gas detectors for high mass WIMPs [37] search in 2012. With a fiducial mass of 34 kg and 225 days of operating time, they set the best limit on the cross section for spin-independent interactions at  $2.0 \times 10^{-9}$  pb for a WIMP mass of 55 GeV. It was then surpassed by LUX, a 370 kg double phase Xenon detector installed in a large water shield, operated in the new SURF (previously Homestake) laboratory in the US. LUX is currently leading the field for masses above 20 GeV, thanks to a run of 85 days with a fiducial mass of 118 kg, setting a limit of  $7.6 \times 10^{-10}$  pb for a WIMP mass of 33 GeV [59]. This data set provides also the best limit for spin dependent WIMPs with pure neutron couplings at all masses [60]. A 300 days run is in progress, allowing an expected factor 5 of improvement in sensitivity.

XENON1t, the successor of XENON100, being installed at the Gran Sasso lab, is going to take data soon. It is expected to overtake LUX in the coming year, if the backgrounds are kept within specifications.

Another liquid Xenon based project, PandaX-I, a double phase detector with pancake geometry, is been operated in the Jinping lab. The latest result obtained with a fiducial mass of 54 kg and a running time of 80 days [61] is competitive with other Xenon experiments only for WIMP masses below 5.5 GeV. PandaX-II, an upgrade of the detector to a mass of 500 kg, is in preparation.

XMASS [30], a single-phase 800 kg Xenon detector (100 kg fiducial mass, allowing a strong self shielding) operated in Japan at the SuperKamiokande site, has seen its detector repaired. Initial commissioning revealed strong radioactive contamination of aluminum pieces close to the photomultipliers. First data taking after refurbishment show a reduction of the differential background rate by a factor 10. Data are being taken. The next step of XMASS is XMASS-1.5 with a 1.5 ton fiducial mass.

The ArDM-1t detector [30], an Argon detector with a total mass of 1.1 t installed at the Canfranc laboratory, has begun operations in the single phase mode. Plans are under way to upgrade the detector with a TPC field cage, allowing to operate it in double phase mode.

DarkSide50, installed in LNGS, is a two phase liquid argon TPC with fiducial mass of 46 kg. The detector is immersed in a spherical vessel containing 30 t of liquid scintillator, which in turn is immersed in a tank containing a kt of pure water. First results from a run of around 30 days [62] have been obtained with natural Argon. After a series of cuts no event was found, leading to the currently best Argon based limit of  $6.1 \times 10^{-8}$  pb for a WIMP mass of 100 GeV. In future they plan is to use Argon from underground sources, which is depleted in the radioactive isotope  $^{39}\text{Ar}$ , reducing the background from this source by a factor of at least 150. If nearly all events recorded in the current run are due to  $^{39}\text{Ar}$  decays, this should allow them to increase the exposure by a similar amount without any background events passing the cuts, so that the sensitivity would scale like the inverse of the exposure. Alternatively they could relax some of the cuts, thereby increasing sensitivity at small WIMP masses.

DEAP-3600 and MiniCLEAN [30], both designed to operate in single phase mode in spherical geometries, are being assembled at SNOLab and will operate 500 kg of Ar/Ne and 3600 kg of Ar, respectively [1]. DEAP-3600 is foreseen to start operation by the end of 2015. The current status of MiniCLEAN, which was expected to undertake liquefaction of argon in the summer of 2014, is unclear.

Candidates for the next generation of multiton Ar and Xe detectors are LZ, XENONnT, DARWIN, DEAP-50T, and DarkSide-20k. Among them, the 15t LZ project has been officially accepted in the US.

The low pressure Time Projection Chamber technique is currently the only convincing way to measure the direction of nuclear recoils and prove the galactic origin of a possible signal [1]. The DRIFT collaboration [30] has operated a  $1\text{ m}^3$  volume detector filled with CS<sub>2</sub> in the UK Boulby mine. Results from a 43 days run with a target mass of 32 g of Fluorine did not show any candidate event but could not probe WIMP models not already excluded by other experiments. The MIMAC collaboration [30] operates a 2.5 l 1000 channel prototype in the Fréjus laboratory, and found first detection of tracks of radon progeny recoils. Other groups developing similar techniques, though with lower sensitivity, are DMTPC in the US and NewAge in Japan.

The two following experiments aim at search for very low mass WIMPs, with mass down to 0.1 GeV. DAMIC [30], using CCDs at SNOLAB, obtained a threshold of around 100 eV. As yet unpublished results from a run with a sensitive mass of 10 g of Silicon provide the best limit in the 1.5-3 GeV mass range. There are plans to increase the sensitive mass to 100 g. The NEWS collaboration [30] exploits an unconventional gas detector, based on a spherical geometry, able to achieve a very low energy threshold, down to a single ionization electron. A 60 cm diameter prototype, SEDINE, is being operated in the Fréjus laboratory. Results from a run using Neon gas should be available soon. A more ambitious project, NEWS-SNO, involving a 1.4 m diameter spherical detector proposed at SNOLAB, hopes to reach sensitivity to WIMP masses down to 0.1 GeV using Hydrogen as target.

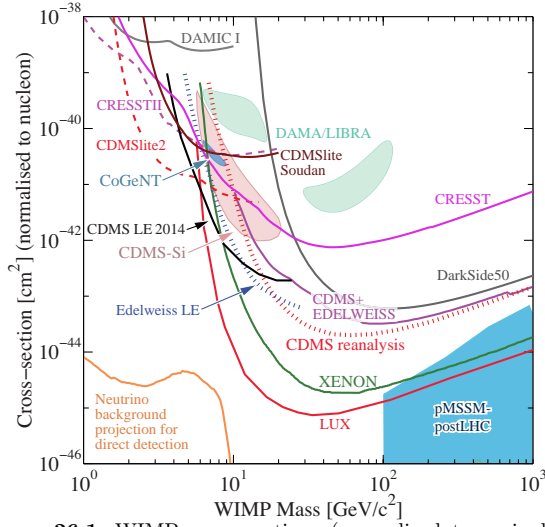
Detectors based on metastable liquids or gels have the advantage of being insensitive to electromagnetic interactions and the drawback of being threshold yes/no detectors. PICO, the merging of the Picasso and COUP collaborations, has operated PICO2L, a bubble chamber type detector filled with 2.9 kg of C<sub>3</sub>F<sub>8</sub> at SNOLAB; a 212 kg-d exposure [63], consisting of runs at different temperatures and hence different thresholds, provided 12 candidates, identified as originating from instrumental imperfection. This led to a limit on the spin dependent proton cross section of  $1 \times 10^{-3}$  pb for a WIMP mass of 30 GeV. This experiment has the best sensitivity worldwide for direct WIMP searches at all masses, but under standard assumptions their limit is weaker than that derived from the bound on WIMP-induced muon neutrinos from the Sun (see below). The PICO60 detector, housing 37 kg of CF<sub>3</sub>I, has been run for more than 12 months but exhibited a large number of anomalous nuclear recoil like events. The collaboration plans to improve the quality of the fluid and to run with C<sub>3</sub>F<sub>8</sub>. The final goal is to build PICO-250L, a ton scale detector.

SIMPLE [30], an experiment using superheated liquid C<sub>2</sub>ClF<sub>5</sub> droplet detectors run at Laboratoire Souterrain de Rustrel, has completed its “phase II”, without bringing better limits than the experiments cited above. The collaboration intends to switch to the bubble chamber technology.

Figures 26.1 and 26.2 illustrate the limits and positive claims for WIMP scattering cross sections, normalized to scattering on a single nucleon, for spin independent and spin dependent couplings, respectively, as functions of WIMP mass. Only the two or three currently best limits are presented. Also shown are constraints from indirect observations (see the next section) and a typical region of a SUSY model after the LHC run-1 results. These figures have been made with the `dmttools` web page [64].

Table 25.1 summarizes the best experimental performances in terms of the upper limit on cross sections for spin independent and spin dependent couplings, at the optimized WIMP mass of each experiment. Also included are some new significant results (using Argon for example).

In summary, the confused situation at low WIMP mass has been cleared up. Many new projects focus on the very low mass range of 0.1-10 GeV. Sensitivities down to  $\sigma_{\chi p}$  of  $10^{-13}$  pb, as needed to probe nearly all of the MSSM parameter space [35] at WIMP masses above 10 GeV and to saturate the limit of the irreducible neutrino-induced background [57], will be reached with Ar and/or Xe detectors of



**Figure 26.1:** WIMP cross sections (normalized to a single nucleon) for spin-independent coupling versus mass. The DAMA/LIBRA [65], CDMS-Si, and CoGeNT enclosed areas are regions of interest from possible signal events. References to the experimental results are given in the text. For context, the blue shaded region shows a scan of the parameter space of the pMSSM, a version of the MSSM with 19 parameters, by the ATLAS collaboration [66], which integrates constraints set by LUX and ATLAS Run 1; the favored region is around  $10^{-10}$  pb and 500 GeV.

multi ton masses, assuming nearly perfect background discrimination capabilities. For WIMP masses below 10 GeV, this cross section limit is set by the solar neutrinos, inducing an irreducible background at an equivalent cross section around  $10^{-9}$  pb, which is accessible with less massive low threshold detectors [30].

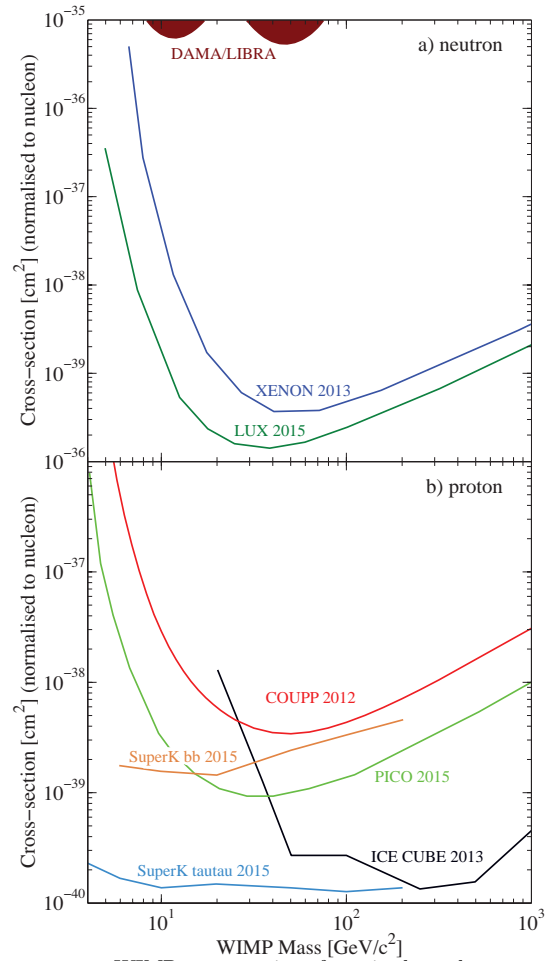
#### 26.2.6. Status and prospects of indirect WIMP searches :

WIMPs can annihilate and their annihilation products can be detected; these include neutrinos, gamma rays, positrons, antiprotons, and antinuclei [1]. These methods are complementary to direct detection and might be able to explore higher masses and different coupling scenarios. “Smoking gun” signals for indirect detection are GeV neutrinos coming from the center of the Sun or Earth, and monoenergetic photons from WIMP annihilation in space.

WIMPs can be slowed down, captured, and trapped in celestial objects like the Earth or the Sun, thus enhancing their density and their probability of annihilation. This is a source of muon neutrinos which can interact in the Earth. Upward going muons can then be detected in large neutrino telescopes such as MACRO, BAKSAN, SuperKamiokande, Baikal, AMANDA, ANTARES, NESTOR, and the large sensitive area IceCube [1]. For standard halo velocity profiles, only the limits from the Sun, which mostly probe spin-dependent couplings, are competitive with direct WIMP search limits.

The best upper limit for WIMP masses up to 200 GeV comes from SuperKamiokande [30]. By including events where the muon is produced inside the detector, in addition to the upgoing events used in earlier analyses, they have been able to extend the sensitivity to the few GeV regime. For example, for WIMPs annihilating into  $b\bar{b}$  pairs, the resulting upper limit on the spin-dependent scattering cross section on protons is about 1.5 (2.3) fb for  $m_\chi = 10$  (50) GeV; for WIMPs annihilating exclusively into  $\tau^+\tau^-$  pairs the bounds are about one order of magnitude stronger [67]. These upper bounds are more than two orders of magnitude below the cross sections required to explain the DAMA signal through spin-dependent scattering on protons.

For heavier WIMPs, giving rise to more energetic muons, the best bounds have been derived from a combination of AMANDA and IceCube40 data (i.e. data using 40 strings of the IceCube detector). For example, for a 1 TeV WIMP annihilating into  $W^+W^-$  the upper



**Figure 26.2:** WIMP cross sections for spin dependent coupling versus mass. (a) interactions with the neutron; (b) interactions with the proton. References to the experimental results are given in the text. Indirect detection results are from SuperKamiokande (annihilation into  $b\bar{b}$  and  $\tau^+\tau^-$  channels) together with IceCube (annihilation into  $W^+W^-$ ); for details see the indirect WIMP searches section below.

bound on the spin-dependent scattering cross section on protons is about 0.25 fb; for WIMPs exclusively annihilating into  $b\bar{b}$  the bound is about 30 times worse [68]. In the future, data including the DeepCore array, which has become part of the completed IceCube detector, will likely dominate this field, possibly except at the very lowest muon energies. However, published bounds from DeepCore in combination with IceCube79 [69] are still weaker than those from SuperKamiokande for relatively soft muons, and are weaker than the combined AMANDA / IceCube40 bound for very energetic muons. These bounds have not changed in the last two years.

WIMP annihilation in the halo can give a continuous spectrum of gamma rays and (at one-loop level) also monoenergetic photon contributions from the  $\gamma\gamma$  and  $\gamma Z$  channels. These channels also allow to search for WIMPs for which direct detection experiments have little sensitivity, e.g., almost pure higgsinos. The size of this signal depends strongly on the halo model, but is expected to be most prominent near the galactic center. The central region of our galaxy hosts a strong TeV point source discovered [70] by the H.E.S.S. Cherenkov telescope [30]. Moreover, Fermi-LAT [30] data revealed a new extended source of GeV photons near the galactic center above and below the galactic plane, the so-called Fermi bubbles [71], as well as several dozen point sources of GeV photons in the inner kpc of our galaxy [72]. These sources are very likely of (mostly) astrophysical origin. The presence of these unexpected backgrounds makes it more difficult to discover WIMPs in this channel.



**Table 26.1:** Summary of performances of the best direct detection experiments, for spin independent and spin dependent couplings. For the “low mass” section, in most cases, there is no minimum in the exclusion curve and a best “typical” WIMP mass cross section point has been chosen.

	Target	Fiducial Mass [kg]	Cross section [pb]	WIMP mass [GeV]	Ref.
<b>Spin independent low mass (&gt;10GeV)</b>					
LUX	Xe	118	$7.6 \times 10^{-10}$	30	[59]
Xenon100	Xe	34	$2.0 \times 10^{-9}$	55	[37]
CDMS/EDW	Ge	12	$2.0 \times 10^{-8}$	100	[55]
DarkSide	Ar	46	$6.1 \times 10^{-8}$	100	[62]
CRESST	CaWO <sub>4</sub> -W	4	$1 \times 10^{-6}$	50	[56]
<b>Spin independent low mass (&lt;10GeV)</b>					
LUX	Xe	118	$1 \times 10^{-8}$	10	[59]
SuperCDMS	Ge LE	$\approx 4.2$	$5 \times 10^{-7}$	10	[52]
SuperCDMS	Ge LE	$\approx 4.2$	$3 \times 10^{-5}$	5	[52]
SuperCDMS	Ge HV	0.6	$3 \times 10^{-4}$	3.3	[53]
CRESST	CaWO <sub>4</sub> -O	0.25	$2 \times 10^{-3}$	2.3	[56]
DAMIC	Si	0.01	$1 \times 10^{-2}$	1.5	[30]
<b>Spin dependent p</b>					
PICO	F	2.9	$1 \times 10^{-3}$	30	[63]
<b>Spin dependent n</b>					
LUX	Xe	118	$3 \times 10^{-4}$	40	[60]

Nevertheless in 2012 a feature was found [73] in public Fermi-LAT data using a predetermined search region around the galactic center, where known point sources had been removed. Within the resolution of the detector this feature could be due to monoenergetic photons with energy  $\sim 130$  GeV. The “local” (in energy and search region) significance of this excess was estimated as 4.6 standard deviations [73], which may have been an over-estimate. In the most recent analysis, based on 5.8 years of data analyzed using the “Pass 8” criteria this feature is no longer visible [74].

Similarly, analyses of publicly available Fermi-LAT data claimed an excess of events in the few GeV range from an extended region around the center of our galaxy, consistent with several WIMP interpretations [75]. A recent, still unpublished, analysis by the Fermi-LAT collaboration [72] indeed found evidence for emission of GeV photons from this region not accounted for by their modelling of astrophysical sources. However, not all of this residual emission can be described by adding a component which is symmetric around the galactic center, as expected for photons from WIMP annihilation or decay. Moreover, the size and spectrum of the fitted “excess” depends strongly on the details of the fits; note that most photons detected from directions around the galactic center actually originate from astrophysical foregrounds, not from the central region, and this foreground is not well understood. Since no error on the estimated total flux from astrophysical sources is given, the statistical evidence for the “excess” cannot be estimated. The collaboration concludes that “a precise physical interpretation of its origin is premature”.

Due to the large astrophysical background near the galactic center, the best bound on WIMPs annihilating into photons in today’s universe comes from a combination of Fermi-LAT observations of dwarf galaxies [76]. It excludes WIMPs annihilating either hadronically or into  $\tau^+\tau^-$  pairs with the standard cross section needed for thermal relics, if the WIMP mass is below  $\sim 100$  GeV; the main assumption is annihilation from an  $S$ -wave initial state. Only slightly weaker limits can be derived from detailed analyses of the CMB by the Planck satellite [77]. The CMB bound assumes otherwise standard cosmology, but also holds if WIMPs dominantly annihilate into light charged leptons.

Antiparticles arise as additional WIMP annihilation products in the halo. To date the best measurements of the antiproton flux come from the PAMELA satellite and the BESS Polar balloon mission [30], and covers kinetic energies between 60 MeV and 180 GeV [78]. The result is in good agreement with secondary production and propagation models. These data exclude WIMP models that attempt to explain the “ $e^\pm$  excesses” (see below) via annihilation into  $W^\pm$  or  $Z^0$  boson pairs; however, largely due to systematic uncertainties they do not significantly constrain conventional WIMP models.

The best measurements of the positron (and electron) flux at energies of tens to hundreds GeV come from AMS02 [79] and PAMELA [80], showing a rather marked rise of the positron fraction between 10 and 200 GeV; the AMS02 data are compatible with a flattening of the positron fraction at the highest energies. While the observed positron spectrum falls within the one order of magnitude span (largely due to differences in the propagation model used) of fluxes predicted by secondary production models [81], the increase of the positron fraction is difficult to reconcile with the rather hard electron spectrum measured by PAMELA [82], if all positrons were due to secondary interactions of cosmic ray particles. Measurements of the total electron+positron energy spectrum by ATIC [83], Fermi-LAT [84] and H.E.S.S. [85] between 100 and 1000 GeV also exceed the predicted purely secondary spectrum, but with very large dispersion of the magnitude of these excesses. These observations can in principle be explained through WIMP annihilation. However, this requires cross sections well above that indicated by Eq. (26.6) for a thermal WIMP. This tension can be resolved only in somewhat baroque WIMP models. Most of these models have by now been excluded by the stringent bounds from Fermi-LAT and from analyses of the CMB on the flux of high energy photons due to WIMP annihilation. This is true also for models trying to explain the leptonic excesses through the decay of WIMPs with lifetime of the order of  $10^{26}$  s. In contrast, viable astrophysical explanations of these excesses introducing new primary sources of electrons and positrons, e.g. pulsars [86] or a nearby supernova that exploded about two million years ago [87], have been suggested. On the other hand, the high quality of the AMS02 data on the positron fraction, which does not show any marked features, allows one to impose stringent bounds on WIMPs with mass below 300 GeV annihilating directly into leptons [88].

Last but not least, an antideuteron signal [1], as potentially observable by AMS02 or PAMELA, could constitute a signal for WIMP annihilation in the halo.

An interesting comparison of respective sensitivities to MSSM parameter space of future direct and various indirect searches has been performed with the DARKSUSY tool [89]. A web-based up-to-date collection of results from direct WIMP searches, theoretical predictions, and sensitivities of future experiments can be found in [64]. Also, the web page [90] allows to make predictions for WIMP signals in various experiments, within a variety of SUSY models and to extract limits from simply parametrised data. Integrated analysis of all data from direct and indirect WIMP detection, and also from LHC experiments should converge to a comprehensive approach, required to fully unravel the mysteries of dark matter.

## References:

1. For details, recent reviews and many more references about particle dark matter, see G. Bertone, *Particle Dark Matter* (Cambridge University Press, 2010).
2. For a brief but delightful history of DM, see V. Trimble, in *Proceedings of the First International Symposium on Sources of Dark Matter in the Universe*, Bel Air, California, 1994, published by World Scientific, Singapore (ed. D.B. Cline). See also the recent review G. Bertone, D. Hooper, and J. Silk, *Phys. Reports* **405**, 279 (2005).
3. E.W. Kolb and M.E. Turner, *The Early Universe*, Addison-Wesley (1990).
4. M. Milgrom, *Can. J. Phys.* **93**, 107 (2015), and references therein.
5. See e.g. L. Bernard, L. Blanchet, *Phys. Rev. D* **91**, 103536(2015).
6. See the section on *Experimental Tests of Gravitational Theory* in this *Review*.



7. See *Cosmological Parameters* in this *Review*.
8. B. Paczynski, *Astrophys. J.* **304**, 1 (1986);  
K. Griest, *Astrophys. J.* **366**, 412 (1991).
9. F. De Paolis *et al.*, *Phys. Rev. Lett.* **74**, 14 (1995).
10. R. Catena and P. Ullio, *JCAP* **1008**, 004 (2010).
11. M. Pato *et al.*, *Phys. Rev.* **D82**, 023531 (2010).
12. N. Bernal *et al.*, *JCAP* **1409**, 004 (2014).
13. J. Bovy and S. Tremaine, *Astrophys. J.* **756**, 89 (2012).
14. K. Kohri, D.H. Lyth, and A. Melchiorri, *JCAP* **0804**, 038 (2008).
15. See *Axions and Other Very Light Bosons* in this *Review*.
16. A. Kusenko, *Phys. Reports* **481**, 1 (2009).
17. For a general introduction to SUSY, see the section devoted in this *Review of Particle Physics*. For a review of SUSY Dark Matter, see G. Jungman, M. Kamionkowski, and K. Griest, *Phys. Reports* **267**, 195 (1996).
18. See *Searches for WIMPs and Other Particles* in this *Review*.
19. T. Moroi and L. Randall, *Nucl. Phys.* **B570**, 455 (2000).
20. R. Allahverdi and M. Drees, *Phys. Rev. Lett.* **89**, 091302 (2002).
21. M. Fujii and T. Yanagida, *Phys. Lett.* **B542**, 80 (2002).
22. J. Hisano, K. Kohri, and M.M. Nojiri, *Phys. Lett.* **B505**, 169 (2001).
23. D.E. Kaplan, M.A. Luty, and K.M. Zurek, *Phys. Rev.* **D79**, 115016 (2009).
24. G. Belanger *et al.*, *Phys. Lett.* **B726**, 773 (2013).
25. J. Kozaczuk and S. Profumo, *Phys. Rev.* **D89**, 095012 (2014).
26. N. Arkani-Hamed *et al.*, *Phys. Rev.* **D79**, 015014 (2009).
27. A. Soffner, [arXiv:1507.02330](https://arxiv.org/abs/1507.02330) [hep-ex], and references therein.
28. Y. Hochberg *et al.*, *Phys. Rev. Lett.* **113**, 171301 (2014).
29. MACHO Collab., C. Alcock *et al.*, *Astrophys. J.* **542**, 257 (2000);  
EROS Collab., *Astron. & Astrophys.* **469**, 387 (2007);  
OGLE Collab., *MNRAS* **416**, 2949 (2011).
30. A very useful collection of web links to the homepages of Dark Matter related conferences, and of experiments searching for WIMP Dark Matter, is the “Dark Matter Portal” at <http://lpsc.in2p3.fr/mayet/dm.php>. See TAUP and IDM conference series sites : <http://www.taup-conference.to.infn.it> and <http://kicp-workshops.uchicago.edu/IDM2012/overview.php>.
31. S.J. Asztalos *et al.*, *Phys. Rev.* **D69**, 011101 (2004);  
S.J. Asztalos *et al.*, *Phys. Rev. Lett.* **104**, 041301 (2010).
32. L.D. Duffy *et al.*, *Phys. Rev.* **D74**, 012006 (2006);  
J. Hoskins *et al.*, *Phys. Rev.* **D84**, 121302 (2011).
33. T.E. Jeltema and S. Profumo, *MNRAS* **450**, 2143 (2015) and references therein..
34. M.C. Smith *et al.*, *MNRAS* **379**, 755 (2007).
35. J. Ellis *et al.*, *Phys. Rev.* **D77**, 065026 (2008).
36. C.E. Aalseth *et al.*, *Phys. Rev. Lett.* **106**, 131301 (2011).
37. XENON100 Collab., E. Aprile *et al.*, *Phys. Rev. Lett.* **109**, 181301 (2012).
38. C.E. Aalseth *et al.*, [arXiv:1401.6234](https://arxiv.org/abs/1401.6234) and [arXiv:1401.3295](https://arxiv.org/abs/1401.3295).
39. J.H. Davis *et al.*, *JCAP* **1408**, 014 (2014).
40. DAMA Collab., R. Bernabei *et al.*, *Phys. Part. Nucl.* **46** (2015) 2, 138-146 (2015).
41. M. Fairbairn and T. Schwetz, *JCAP* **0901**, 037 (2009).
42. DAMA Collab., R. Bernabei *et al.*, *Eur. Phys. J.* **C56**, 333 (2008).
43. DAMA Collab., R. Bernabei *et al.*, *Eur. Phys. J.* **C67**, 39 (2010).
44. H.S. Lee *et al.*, *JHEP* **08**, 093 (2015).
45. DM-ICE Collab., J. Cherwinka *et al.*, [arXiv:1509.02486](https://arxiv.org/abs/1509.02486).
46. N. Bozorgnia and T. Schwetz, *JCAP* **1412**, 015 (2014).
47. S.C. Kim *et al.*, *Phys. Rev. Lett.* **108**, 181301 (2012).
48. H. S. Lee *et al.*, *Phys. Rev.* **D90**, 052006 (2014).
49. J. Klinger and V.A. Kudryavtsev, *Phys. Rev. Lett.* **114**, 151301 (2015).
50. Q. Wallemacq and J.R. Cudell, *JCAP* **1502**, 011 (2015);  
R. Foot, *Phys. Rev.* **D90**, 121302 (2014).
51. SuperCDMS Collab., *Phys. Rev.* **D92**, 072003 (2015).
52. SuperCDMS Collab., *Phys. Rev. Lett.* **112**, 241302 (2014).
53. SuperCDMS Collab., *Phys. Rev. Lett.* **112**, 041302 (2014);  
SuperCDMS Collab., [arXiv:1509.02448](https://arxiv.org/abs/1509.02448).
54. EDELWEISS Collab., [arXiv:1504.00820](https://arxiv.org/abs/1504.00820).
55. EDELWEISS and CDMS Collab., Z. Ahmed *et al.*, *Phys. Rev.* **84**, 011102 (2011).
56. CRESST-II Collab., G. Angloher *et al.*, *Eur. Phys. J.* **C74**, 3184 (2014); CRESST-II Collab., G. Angloher *et al.*, [arXiv:1509.01515](https://arxiv.org/abs/1509.01515).
57. J. Billard, L. Strigari, E. Figueroa-Feliciano, *Phys. Rev.* **D89**, 023524 (2014).
58. SNOWMASS Working Group Report: FERMILAB-CONF-13-688-AE,  
[arXiv:1310.8327](https://arxiv.org/abs/1310.8327).
59. LUX Collab., D.S. Akerib *et al.*, *Phys. Rev. Lett.* **112**, 091303 (2014).
60. C. Savage *et al.*, *Phys. Rev.* **D92**, 103519 (2015).
61. PandaX Collab., *Phys. Rev.* **D92**, 052004 (2015).
62. DarkSide Collab., *Phys. Lett.* **B743**, 456 (2015).
63. PICO Collab. *et al.*, *Phys. Rev. Lett.* **114**, 231302 (2015).
64. DMTOOLS site : <http://dmtools.brown.edu:8080/>.
65. C. Savage *et al.*, *JCAP* **0904**, 010, (2009).
66. M. Cahill-Rowley *et al.*, [arXiv:1305.6921v2](https://arxiv.org/abs/1305.6921v2) and ATLAS Collab. *JHEP* **1510**, 134 (2015).
67. SuperKamiokande Collab., K. Choi *et al.*, *Phys. Rev. Lett.* **114**, 141301 (2015).
68. IceCube Collab., R. Abbasi *et al.*, *Phys. Rev.* **D85**, 042002 (2012).
69. IceCube Collab., M.G. Aartsen *et al.*, *Phys. Rev. Lett.* **110**, 131302 (2013).
70. H.E.S.S. Collab., F. Aharonian *et al.*, *Astron. & Astrophys.* **503**, 817 (2009);  
H.E.S.S. Collab., F. Acero *et al.*, *MNRAS* **402**, 1877 (2010).
71. M. Su, T.R. Slatyer, and D.P. Finkbeiner, *Astrophys. J.* **724**, 1044 (2010).
72. T.A. Porter and S. Murgia for the Fermi-LAT Collab., [arXiv:1507.04688](https://arxiv.org/abs/1507.04688).
73. C. Weniger, *JCAP* **1208**, 007 (2012).
74. Fermi-LAT Collab., M. Ackermann *et al.*, *Phys. Rev.* **D91**, 122002 (2015).
75. D. Hooper and L. Goodenough, *Phys. Lett.* **B697**, 412 (2011);  
T. Daylan *et al.*, [arXiv:1402.6703](https://arxiv.org/abs/1402.6703).
76. Fermi-LAT Collab., M. Ackermann *et al.*, [arXiv:1503.02641](https://arxiv.org/abs/1503.02641).
77. G. Steigman, *Phys. Rev.* **D91**, 083538 (2015).
78. PAMELA Collab., O. Adriani *et al.*, *Phys. Rev. Lett.* **105**, 121101 (2010);  
BESS Polar Collab., K. Abe *et al.*, *Phys. Rev. Lett.* **108**, 051102 (2012).
79. AMS02 Collab., M. Aguilar *et al.*, *Phys. Rev. Lett.* **113**, 121101 (2014) and *Phys. Rev. Lett.* **113**, 121102 (2014).
80. PAMELA Collab., O. Adriani *et al.*, *Phys. Rev. Lett.* **111**, 081102 (2013).
81. T. Delahaye *et al.*, *Astron. & Astrophys.* **501**, 821 (2009).
82. PAMELA Collab., O. Adriani *et al.*, *Phys. Rev. Lett.* **106**, 201101 (2011).
83. ATIC Collab, J. Chang *et al.*, *Nature* **456**, 362 (2008).
84. FERMI/LAT Collab, A.A. Abdo *et al.*, *Phys. Rev. Lett.* **102**, 181101 (2009).
85. H.E.S.S. Collab, F. Aharonian *et al.*, *Astron. & Astrophys.* **508**, 561 (2009).
86. M. Cirelli, *Pramana* **79**, 1021 (2012);  
S. Profumo, *Central Eur. J. Phys.* **10**, 1 (2011).
87. M. Kachelriess, A. Neronov, D.V. Semikoz, *Phys. Rev. Lett.* **115**, 181103 (2015).
88. L. Bergstrom *et al.*, *Phys. Rev. Lett.* **111**, 171101 (2013).
89. DARKSUSY site: <http://www.physto.se/edsjo/darksusy/>.
90. ILIAS web page: <http://pisrv0.pit.physik.uni-tuebingen.de/darkmatter/>.

## 27. DARK ENERGY

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### 27.1. Repulsive Gravity and Cosmic Acceleration

In the first modern cosmological model, Einstein [1] modified his field equation of General Relativity (GR), introducing a “cosmological term” that enabled a solution with time-independent, spatially homogeneous matter density  $\rho_m$  and constant positive space curvature. Although Einstein did not frame it this way, one can view the “cosmological constant”  $\Lambda$  as representing a constant energy density of the vacuum [2], whose repulsive gravitational effect balances the attractive gravity of matter and thereby allows a static solution. After the development of dynamic cosmological models [3,4] and the discovery of cosmic expansion [5], the cosmological term appeared unnecessary, and Einstein and de Sitter [6] advocated adopting an expanding, homogeneous and isotropic, spatially flat, matter-dominated universe as the default cosmology until observations dictated otherwise. Such a model has matter density equal to the critical density,  $\Omega_m \equiv \rho_m/\rho_c = 1$ , and negligible contribution from other energy components [7].

By the mid-1990s, the Einstein-de Sitter model was showing numerous cracks, under the combined onslaught of data from the cosmic microwave background (CMB), large-scale galaxy clustering, and direct estimates of the matter density, the expansion rate ( $H_0$ ), and the age of the Universe. As noted in a number of papers from this time, introducing a cosmological constant offered a potential resolution of many of these tensions, yielding the most empirically successful version of the inflationary cold dark matter scenario. In the late 1990s, supernova surveys by two independent teams provided direct evidence for accelerating cosmic expansion [8,9], establishing the cosmological constant model (with  $\Omega_m \approx 0.3$ ,  $\Omega_\Lambda \approx 0.7$ ) as the preferred alternative to the  $\Omega_m = 1$  scenario. Shortly thereafter, CMB evidence for a spatially flat universe [10,11], and thus for  $\Omega_{\text{tot}} \approx 1$ , cemented the case for cosmic acceleration by firmly eliminating the free-expansion alternative with  $\Omega_m \ll 1$  and  $\Omega_\Lambda = 0$ . Today, the accelerating universe is well established by multiple lines of independent evidence from a tight web of precise cosmological measurements.

As discussed in the Big Bang Cosmology article of this *Review* (Sec. 22), the scale factor  $R(t)$  of a homogeneous and isotropic universe governed by GR grows at an accelerating rate if the pressure  $p < -\frac{1}{3}\rho$  (in  $c = 1$  units). A cosmological constant has  $\rho_\Lambda = \text{constant}$  and pressure  $p_\Lambda = -\rho_\Lambda$  (see Eq. 22.10), so it will drive acceleration if it dominates the total energy density. However, acceleration could arise from a more general form of “dark energy” that has negative pressure, typically specified in terms of the equation-of-state-parameter  $w = p/\rho$  ( $= -1$  for a cosmological constant). Furthermore, the conclusion that acceleration requires a new energy component beyond matter and radiation relies on the assumption that GR is the correct description of gravity on cosmological scales. The title of this article follows the common but inexact usage of “dark energy” as a catch-all term for the origin of cosmic acceleration, regardless of whether it arises from a new form of energy or a modification of GR. Our account here draws on the much longer review of cosmic acceleration by Ref. [12], which provides background explanation and extensive literature references for most of the points in this article, but is less up to date in its description of current empirical constraints.

Below we will use the abbreviation  $\Lambda\text{CDM}$  to refer to a model with cold dark matter, a cosmological constant, inflationary initial conditions, standard radiation and neutrino content, and a flat universe with  $\Omega_{\text{tot}} = 1$  (though we will sometimes describe this model as “flat  $\Lambda\text{CDM}$ ” to emphasize this last restriction). We will use  $w\text{CDM}$  to denote a model with the same assumptions but a free, constant value of  $w$ . Models with the prefix “o” (*e.g.*,  $ow\text{CDM}$ ) allow non-zero space curvature.

### 27.2. Theories of Cosmic Acceleration

#### 27.2.1. Dark Energy or Modified Gravity? :

A cosmological constant is the mathematically simplest, and perhaps the physically simplest, theoretical explanation for the accelerating universe. The problem is explaining its unnaturally small magnitude, as discussed in Sec. 22.4.7 of this *Review*. An alternative (which still requires finding a way to make the cosmological constant zero or at least negligibly small) is that the accelerating cosmic expansion is driven by a new form of energy such as a scalar field [13] with potential  $V(\phi)$ . The energy density and pressure of the field  $\phi(\mathbf{x})$  take the same forms as for inflationary scalar fields, given in Eq. (22.52) of the Big Bang Cosmology article. In the limit that  $\frac{1}{2}\dot{\phi}^2 \ll |V(\phi)|$ , the scalar field acts like a cosmological constant, with  $p_\phi \approx -\rho_\phi$ . In this scenario, today’s cosmic acceleration is closely akin to the epoch of inflation, but with radically different energy and timescale.

More generally, the value of  $w = p_\phi/\rho_\phi$  in scalar field models evolves with time in a way that depends on  $V(\phi)$  and on the initial conditions  $(\phi_i, \dot{\phi}_i)$ ; some forms of  $V(\phi)$  have attractor solutions in which the late-time behavior is insensitive to initial values. Many forms of time evolution are possible, including ones where  $w$  is approximately constant and broad classes where  $w$  “freezes” towards or “thaws” away from  $w = -1$ , with the transition occurring when the field comes to dominate the total energy budget. If  $\rho_\phi$  is even approximately constant, then it becomes dynamically insignificant at high redshift, because the matter density scales as  $\rho_m \propto (1+z)^3$ . “Early dark energy” models are ones in which  $\rho_\phi$  is a small but not negligible fraction (*e.g.*, a few percent) of the total energy throughout the matter- and radiation-dominated eras, tracking the dominant component before itself coming to dominate at low redshift.

Instead of introducing a new energy component, one can attempt to modify gravity in a way that leads to accelerated expansion [14]. One option is to replace the Ricci scalar  $\mathcal{R}$  with a function  $\mathcal{R} + f(\mathcal{R})$  in the gravitational action [15]. Other changes can be more radical, such as introducing extra dimensions and allowing gravitons to “leak” off the brane that represents the observable universe (the “DGP” model [16]). The DGP example has inspired a more general class of “galileon” and massive gravity models. Constructing viable modified gravity models is challenging, in part because it is easy to introduce theoretical inconsistencies (such as “ghost” fields with negative kinetic energy), but above all because GR is a theory with many high-precision empirical successes on solar system scales [17]. Modified gravity models typically invoke screening mechanisms that force model predictions to approach those of GR in regions of high density or strong gravitational potential. Screening offers potentially distinctive signatures, as the strength of gravity (*i.e.*, the effective value of  $G_N$ ) can vary by order unity in environments with different gravitational potentials.

More generally, one can search for signatures of modified gravity by comparing the history of cosmic structure growth to the history of cosmic expansion. Within GR, these two are linked by a consistency relation, as described below (Eq. (27.2)). Modifying gravity can change the predicted rate of structure growth, and it can make the growth rate dependent on scale or environment. In some circumstances, modifying gravity alters the combinations of potentials responsible for gravitational lensing and the dynamics of non-relativistic tracers (such as galaxies or stars) in different ways (see Sec. 22.4.7 in this *Review*), leading to order unity mismatches between the masses of objects inferred from lensing and those inferred from dynamics in unscreened environments.

At present there are no fully realized and empirically viable modified gravity theories that explain the observed level of cosmic acceleration. The constraints on  $f(\mathcal{R})$  models now force them so close to GR that they cannot produce acceleration without introducing a separate dark energy component [18]. The DGP model is empirically ruled out by several tests, including the expansion history, the integrated Sachs-Wolfe effect, and redshift-space distortion measurements of the structure growth rate [19]. The elimination of these models should be considered an important success of the program to empirically test theories of cosmic acceleration. However, it is worth recalling that

there was no fully realized gravitational explanation for the precession of Mercury's orbit prior to the completion of GR in 1915, and the fact that no complete and viable modified gravity theory exists today does not mean that one will not arise in the future. In the meantime, we can continue empirical investigations that can tighten restrictions on such theories or perhaps point towards the gravitational sector as the origin of accelerating expansion.

### 27.2.2. Expansion History and Growth of Structure :

The main line of empirical attack on dark energy is to measure the history of cosmic expansion and the history of matter clustering with the greatest achievable precision over a wide range of redshift. Within GR, the expansion rate  $H(z)$  is governed by the Friedmann equation (see the articles on Big Bang Cosmology and Cosmological Parameters—Secs. 22 and 25 in this *Review*). For dark energy with an equation of state  $w(z)$ , the cosmological constant contribution to the expansion,  $\Omega_\Lambda$ , is replaced by a redshift-dependent contribution. The evolution of the dark energy density follows from Eq. (22.10),

$$\Omega_{\text{de}} \frac{\rho_{\text{de}}(z)}{\rho_{\text{de}}(z=0)} = \Omega_{\text{de}} \exp \left[ 3 \int_0^z [1 + w(z')] \frac{dz'}{1+z'} \right] = \Omega_{\text{de}} (1+z)^{3(1+w)}, \quad (27.1)$$

where the second equality holds for constant  $w$ . If  $\Omega_m$ ,  $\Omega_r$ , and the present value of  $\Omega_{\text{tot}}$  are known, then measuring  $H(z)$  pins down  $w(z)$ . (Note that  $\Omega_{\text{de}}$  is the same quantity denoted  $\Omega_v$  in Sec. 22, but we have adopted the  $\text{de}$  subscript to avoid implying that dark energy is necessarily a vacuum effect.)

While some observations can probe  $H(z)$  directly, others measure the distance-redshift relation. The basic relations between angular diameter distance or luminosity distance and  $H(z)$  are given in Ch. 22—and these are generally unaltered in time-dependent dark energy or modified gravity models. For convenience, in later sections, we will sometimes refer to the comoving angular distance,  $D_{A,c}(z) = (1+z)D_A(z)$ .

In GR-based linear perturbation theory, the density contrast  $\delta(\mathbf{x}, t) \equiv \rho(\mathbf{x}, t)/\bar{\rho}(t) - 1$  of pressureless matter grows in proportion to the linear growth function  $G(t)$  (not to be confused with the gravitational constant  $G_N$ ), which follows the differential equation

$$\ddot{G} + 2H(z)\dot{G} - \frac{3}{2}\Omega_m H_0^2 (1+z)^3 G = 0. \quad (27.2)$$

To a good approximation, the logarithmic derivative of  $G(z)$  is

$$f(z) \equiv -\frac{d \ln G}{d \ln(1+z)} \approx \left[ \Omega_m (1+z)^3 \frac{H_0^2}{H^2(z)} \right]^\gamma, \quad (27.3)$$

where  $\gamma \approx 0.55$  for relevant values of cosmological parameters [20]. In an  $\Omega_m = 1$  universe,  $G(z) \propto (1+z)^{-1}$ , but growth slows when  $\Omega_m$  drops significantly below unity. One can integrate Eq. (27.3) to get an approximate integral relation between  $G(z)$  and  $H(z)$ , but the full (numerical) solution to Eq. (27.2) should be used for precision calculations. Even in the non-linear regime, the amplitude of clustering is determined mainly by  $G(z)$ , so observations of non-linear structure can be used to infer the linear  $G(z)$ , provided one has good theoretical modeling to relate the two.

In modified gravity models the growth rate of gravitational clustering may differ from the GR prediction. A general strategy to test modified gravity, therefore, is to measure both the expansion history and the growth history to see whether they yield consistent results for  $H(z)$  or  $w(z)$ .

### 27.2.3. Parameters :

Constraining a general history of  $w(z)$  is nearly impossible, because the dark energy density, which affects  $H(z)$ , is given by an integral over  $w(z)$ , and distances and the growth factor involve a further integration over functions of  $H(z)$ . Oscillations in  $w(z)$  over a range  $\Delta z/(1+z) \ll 1$  are therefore extremely difficult to constrain. It has become conventional to phrase constraints or projected constraints on  $w(z)$  in terms of a linear evolution model,

$$w(a) = w_0 + w_a(1-a) = w_p + w_a(a_p - a), \quad (27.4)$$

where  $a \equiv (1+z)^{-1}$ ,  $w_0$  is the value of  $w$  at  $z = 0$ , and  $w_p$  is the value of  $w$  at a “pivot” redshift  $z_p \equiv a_p^{-1} - 1$ , where it is best constrained by a given set of experiments. For typical data combinations,  $z_p \approx 0.5$ . This simple parameterization can provide a good approximation to the predictions of many physically motivated models for observables measured with percent-level precision. A widely used “Figure of Merit” (FoM) for dark energy experiments [21] is the projected combination of errors  $[\sigma(w_p)\sigma(w_a)]^{-1}$ . Ambitious future experiments with 0.1–0.3% precision on observables can constrain richer descriptions of  $w(z)$ , which can be characterized by principal components.

There has been less convergence on a standard parameterization for describing modified gravity theories. Deviations from the GR-predicted growth rate can be described by a deviation  $\Delta\gamma$  in the index of Eq. (27.3), together with an overall multiplicative offset relative to the  $G(z)$  expected from extrapolating the CMB-measured fluctuation amplitude to low redshift. However, these two parameters may not accurately capture the growth predictions of all physically interesting models. Another important parameter to constrain is the ratio of the gravitational potentials governing space curvature and the acceleration of non-relativistic test particles. The possible phenomenology of modified gravity models is rich, which enables many consistency tests but complicates the task of constructing parameterized descriptions.

The more general set of cosmological parameters is discussed elsewhere in this *Review* (Sec. 25), but here we highlight a few that are particularly important to the dark energy discussion:

- The dimensionless Hubble parameter  $h \equiv H_0/100 \text{ km s}^{-1} \text{ Mpc}^{-1}$  determines the present day value of the critical density and the overall scaling of distances inferred from redshifts.
- $\Omega_m$  and  $\Omega_{\text{tot}}$  affect the expansion history and the distance-redshift relation.
- The sound horizon  $r_s = \int_0^{t_{\text{rec}}} c_s(t) dt/a(t)$ , the comoving distance that pressure waves can propagate between  $t = 0$  and recombination, determines the physical scale of the acoustic peaks in the CMB and the baryon acoustic oscillation (BAO) feature in low redshift matter clustering [22].
- The amplitude of matter fluctuations, conventionally represented by the quantity  $\sigma_8(z)$ , scales the overall amplitude of growth measures such as weak lensing or redshift-space distortions (discussed in the next section).

Specifically,  $\sigma_8(z)$  refers to the rms fluctuation of the matter overdensity  $\rho/\bar{\rho}$  in spheres of radius  $8 h^{-1} \text{ Mpc}$ , computed from the linear theory matter power spectrum at redshift  $z$ , and  $\sigma_8$  on its own refers to the value at  $z = 0$  (just like our convention for  $\Omega_m$ ).

While discussions of dark energy are frequently phrased in terms of values and errors on quantities like  $w_p$ ,  $w_a$ ,  $\Delta\gamma$ , and  $\Omega_{\text{tot}}$ , parameter precision is the means to an end, not an end in itself. The underlying goal of empirical studies of cosmic acceleration is to address two physically profound questions:

1. Does acceleration arise from a breakdown of GR on cosmological scales or from a new energy component that exerts repulsive gravity within GR?
2. If acceleration is caused by a new energy component, is its energy density constant in space and time, as expected for a fundamental vacuum energy, or does it show variations that indicate a dynamical field?

Substantial progress towards answering these questions, in particular any definitive rejection of the cosmological constant “null hypothesis,” would be a major breakthrough in cosmology and fundamental physics.

## 27.3. Observational Probes

We briefly summarize the observational probes that play the greatest role in current constraints on dark energy. Further discussion can be found in other articles of this *Review*, in particular Secs. 25 (Cosmological Parameters) and 28 (The Cosmic Microwave Background), and in Ref. [12], which provides extensive references to background literature. Recent observational results from these methods are discussed in 27.4.

### 27.3.1. Methods, Sensitivity, Systematics :

*Cosmic Microwave Background Anisotropies:* Although CMB anisotropies provide limited information about dark energy on their own, CMB constraints on the geometry, matter content, and radiation content of the Universe play a critical role in dark energy studies when combined with low redshift probes. In particular, CMB data supply measurements of  $\theta_s = r_s/D_{A,c}(z_{\text{rec}})$ , the angular size of the sound horizon at recombination, from the angular location of the acoustic peaks, measurements of  $\Omega_m h^2$  and  $\Omega_b h^2$  from the heights of the peaks, and normalization of the amplitude of matter fluctuations at  $z_{\text{rec}}$  from the amplitude of the CMB fluctuations themselves. *Planck* data yield a 0.4% determination of  $r_s$ , which scales as  $(\Omega_m h^2)^{-0.25}$  for cosmologies with standard matter and radiation content. The uncertainty in the matter fluctuation amplitude is 1–2%. Improvements in the measurement of the electron scattering optical depth  $\tau$ , with future analyses of *Planck* polarization maps, would reduce this uncertainty further. Secondary anisotropies, including the Integrated Sachs-Wolfe effect and the Sunyaev-Zeldovich (SZ, [23]) effect, provide additional information about dark energy by constraining low-redshift structure growth.

*Type Ia Supernovae:* Type Ia supernovae, produced by the thermonuclear explosions of white dwarfs, exhibit 10–15% scatter in peak luminosity after correction for light curve duration (the time to rise and fall) and color (which is a diagnostic of dust extinction). Since the peak luminosity is not known *a priori*, supernova surveys constrain ratios of luminosity distances at different redshifts. If one is comparing a high redshift sample to a local calibrator sample measured with much higher precision (and distances inferred from Hubble's law), then one essentially measures the luminosity distance in  $h^{-1}\text{Mpc}$ , constraining the combination  $hD_L(z)$ . With distance uncertainties of 5–8% per well observed supernova, a sample of around 100 SNe is sufficient to achieve sub-percent statistical precision. The 1–2% systematic uncertainties in current samples are dominated by uncertainties associated with photometric calibration and dust extinction corrections plus the observed dependence of luminosity on host galaxy properties. Another potential systematic is redshift evolution of the supernova population itself, which can be tested by analyzing subsamples grouped by spectral properties or host galaxy properties to confirm that they yield consistent results.

*Baryon Acoustic Oscillations (BAO):* Pressure waves that propagate in the pre-recombination photon-baryon fluid imprint a characteristic scale in the clustering of matter and galaxies, which appears in the galaxy correlation function as a localized peak at the sound horizon scale  $r_s$ , or in the power spectrum as a series of oscillations. Since observed galaxy coordinates consist of angles and redshifts, measuring this “standard ruler” scale in a galaxy redshift survey determines the angular diameter distance  $D_A(z)$  and the expansion rate  $H(z)$ , which convert coordinate separations to comoving distances. Errors on the two quantities are correlated, and in existing galaxy surveys the best determined combination is approximately  $D_V(z) = [czD_{A,c}^2(z)/H(z)]^{1/3}$ . As an approximate rule of thumb, a survey that fully samples structures at redshift  $z$  over a comoving volume  $V$ , and is therefore limited by cosmic variance rather than shot noise, measures  $D_{A,c}(z)$  with a fractional error of  $0.005(V/10\text{Gpc}^3)^{-1/2}$  and  $H(z)$  with a fractional error 1.6–1.8 times higher. The most precise BAO measurements to date come from large galaxy redshift surveys probing  $z < 0.8$ , and these will be extended to higher redshifts by future projects. At redshifts  $z > 2$ , BAO can also be measured in the Lyman- $\alpha$  forest of intergalactic hydrogen absorption towards background quasars, where the fluctuating absorption pattern provides tens or hundreds of samples of the density field along each quasar sightline. For Lyman- $\alpha$  forest BAO, the best measured parameter combination is more heavily weighted towards  $H(z)$  because of strong redshift-space distortions that enhance clustering in the line-of-sight direction. Radio intensity mapping, which maps large scale structure in redshifted 21cm hydrogen emission without resolving individual galaxies, offers a potentially promising route to measuring BAO over large volumes at relatively low cost, but the technique is still under development. Photometric redshifts in optical imaging surveys can be used to measure BAO in the angular

direction, though the typical distance precision is a factor of 3–4 lower compared to a well sampled spectroscopic survey of the same area, and angular BAO measurements do not directly constrain  $H(z)$ . BAO distance measurements complement SN distance measurements by providing absolute rather than relative distances (with precise calibration of  $r_s$  from the CMB) and by having greater achievable precision at high redshift thanks to the increasing comoving volume available. Theoretical modeling suggests that BAO measurements from even the largest feasible redshift surveys will be limited by statistical rather than systematic uncertainties.

*Weak Gravitational Lensing:* Gravitational light bending by a clustered distribution of matter shears the shapes of higher redshift background galaxies in a spatially coherent manner, producing a correlated pattern of apparent ellipticities. By studying the weak lensing signal for source galaxies binned by photometric redshift (estimated from broad-band colors), one can probe the history of structure growth. For a specified expansion history, the predicted signal scales approximately as  $\sigma_8 \Omega_m^\alpha$ , with  $\alpha \approx 0.3\text{--}0.5$ . The predicted signal also depends on the distance-redshift relation, so weak lensing becomes more powerful in concert with SN or BAO measurements that can pin this relation down independently. The most challenging systematics are shape measurement biases, biases in the distribution of photometric redshifts, and intrinsic alignments of galaxy orientations that could contaminate the lensing-induced signal. Predicting the large-scale weak lensing signal is straightforward in principle, but the number of independent modes on large scales is small, and the inferences are therefore dominated by sample variance. Exploiting small-scale measurements, for tighter constraints, requires modeling the effects of complex physical processes such as star formation and feedback on the matter power spectrum. Strong gravitational lensing can also provide constraints on dark energy, either through time delay measurements that probe the absolute distance scale, or through measurements of multiple-redshift lenses that constrain distance ratios. The primary uncertainty for strong lensing constraints is modeling the mass distribution of the lens systems.

*Clusters of Galaxies:* Like weak lensing, the abundance of massive dark matter halos probes structure growth by constraining  $\sigma_8 \Omega_m^\alpha$ , where  $\alpha \approx 0.3\text{--}0.5$ . These halos can be identified as dense concentrations of galaxies or through the signatures of hot ( $10^7\text{--}10^8\text{K}$ ) gas in X-ray emission or SZ distortion of the CMB. The critical challenge in cluster cosmology is calibrating the relation  $P(M_{\text{halo}}|O)$  between the halo mass as predicted from theory and the observable  $O$  used for cluster identification. Measuring the stacked weak lensing signal from clusters has emerged as a promising approach to achieve percent-level accuracy in calibration of the mean relation, which is required for clusters to remain competitive with other growth probes. This method requires accurate modeling of completeness and contamination of cluster catalogs, projection effects on cluster selection and weak lensing measurements, and possible baryonic physics effects on the mass distribution within clusters.

*Redshift-Space Distortions (RSD) and the Alcock-Paczynski (AP) Effect:* Redshift-space distortions of galaxy clustering, induced by peculiar motions, probe structure growth by constraining the parameter combination  $f(z)\sigma_8(z)$ , where  $f(z)$  is the growth rate defined by Eq. (27.3). Uncertainties in theoretical modeling of non-linear gravitational evolution and the non-linear bias between the galaxy and matter distributions currently limit application of the method to large scales (comoving separations  $r \gtrsim 10 h^{-1}\text{Mpc}$  or wavenumbers  $k \lesssim 0.2 h\text{Mpc}^{-1}$ ). A second source of anisotropy arises if one adopts the wrong cosmological metric to convert angles and redshifts into comoving separations, a phenomenon known as the Alcock-Paczynski effect [24]. Demanding isotropy of clustering at redshift  $z$  constrains the parameter combination  $H(z)D_A(z)$ . The main challenge for the AP method is correcting for the anisotropy induced by peculiar velocity RSD.

*Direct Determination of  $H_0$ :* The value of  $H_0$  sets the current value of the critical density  $\rho_c = 3H_0^2/8\pi G_N$ , and combination with CMB measurements provides a long lever arm for constraining the evolution of dark energy. The challenge in direct  $H_0$  measurements is establishing distances to galaxies that are “in the Hubble flow,”

**Table 27.2:** A selection of major dark energy experiments, based on Ref. [25]. Abbreviations in the “Data” column refer to optical (Opt) or near-infrared (NIR) imaging (I) or spectroscopy (S). For spectroscopic experiments, the “Spec- $z$ ” column lists the primary redshift range for galaxies (gals), quasars (QSOs), or the Lyman- $\alpha$  forest (Ly $\alpha$ F). Abbreviations in the “Methods” column are weak lensing (WL), clusters (CL), supernovae (SN), baryon acoustic oscillations (BAO), and redshift-space distortions (RSD).

Project	Dates	Area/deg <sup>2</sup>	Data	Spec- $z$ Range	Methods
BOSS	2008-2014	10,000	Opt-S	0.3 – 0.7 (gals) 2 – 3.5 (Ly $\alpha$ F)	BAO/RSD
DES	2013-2018	5000	Opt-I	—	WL/CL SN/BAO
eBOSS	2014-2020	7500	Opt-S	0.6 – 2.0 (gal/QSO) 2 – 3.5 (Ly $\alpha$ F)	BAO/RSD
SuMIRE	2014-2024	1500	Opt-I	—	WL/CL
			Opt/NIR-S	0.8 – 2.4 (gals)	BAO/RSD
HETDEX	2014-2019	300	Opt-S	1.9 < $z$ < 3.5 (gals)	BAO/RSD
DESI	2019-2024	14,000	Opt-S	0 – 1.7 (gals) 2 – 3.5 (Ly $\alpha$ F)	BAO/RSD
LSST	2020-2030	20,000	Opt-I	—	WL/CL SN/BAO
<i>Euclid</i>	2020-2026	15,000	Opt-I	—	WL/CL
			NIR-S	0.7 – 2.2 (gals)	BAO/RSD
<i>WFIRST</i>	2024-2030	2200	NIR-I	—	WL/CL/SN
			NIR-S	1.0 – 3.0 (gals)	BAO/RSD

*i.e.*, far enough away that their peculiar velocities are small compared to the expansion velocity  $v = H_0 d$ . This can be done by building a ladder of distance indicators tied to stellar parallax on its lowest rung, or by using gravitational lens time delays or geometrical measurements of maser data to circumvent this ladder.

### 27.3.2. Dark Energy Experiments :

Most observational applications of these methods now take place in the context of large cosmological surveys, for which constraining dark energy and modified gravity theories is a central objective. Table 27.2 lists a selection of current and planned dark energy experiments, taken from the Snowmass 2013 Dark Energy Facilities review [25], which focused on projects in which the U.S. has either a leading role or significant participation. References and links to further information about these projects can be found in Ref. [25].

Beginning our discussion with imaging surveys, the Dark Energy Survey (DES) will cover 1/8 of the sky to a depth roughly 2 magnitudes deeper than the Sloan Digital Sky Survey (SDSS), enabling weak lensing measurements with unprecedented statistical precision, cluster measurements calibrated by weak lensing, and angular BAO measurements based on photometric redshifts. With repeat imaging over a smaller area, DES will identify thousands of Type Ia SNe, which together with spectroscopic follow-up data will enable significant improvements on the current state-of-the-art for supernova (SN) cosmology. The Hyper-Suprime Camera (HSC) on the Subaru 8.2-meter telescope will carry out a similar type of optical imaging survey, probing a smaller area than DES but to greater depth. This survey is one component of the Subaru Measurement of Images and Redshifts (SuMIRE) project. Beginning in the early 2020s, the dedicated Large Synoptic Survey Telescope (LSST) will scan the southern sky to SDSS-like depth every four nights. LSST imaging co-added over its decade-long primary survey will reach extraordinary depth, enabling weak lensing, cluster, and photometric BAO studies from billions of galaxies. LSST time-domain monitoring will identify and measure light curves for thousands of Type Ia SNe per year.

Turning to spectroscopic surveys, the Baryon Oscillation Spectroscopic Survey (BOSS) and its successor eBOSS use fiber-fed optical spectrographs to map the redshift-space distributions of millions of galaxies and quasars. These 3-dimensional maps enable BAO and RSD measurements, and Lyman- $\alpha$  forest spectra of high-redshift quasars extend these measurements to redshifts  $z > 2$ . The Hobby-Eberly Telescope Dark Energy Experiment (HETDEX) uses integral field spectrographs to detect Lyman- $\alpha$  emission-line galaxies at  $z \approx 1.9 - 3.5$ , probing a small sky area but a substantial comoving volume. The Dark Energy Spectroscopic Instrument (DESI) follows a strategy similar to BOSS/eBOSS but on a much grander scale, using a larger telescope (4-meter vs. 2.5-meter) and a much higher fiber multiplex (5000 vs. 1000) to survey an order-of-magnitude more galaxies. A new Prime Focus Spectrograph (PFS) for the Subaru telescope will enable the spectroscopic component of SuMIRE, with the large telescope aperture and wavelength sensitivity that extends to the near-infrared (NIR) allowing it to probe a higher redshift galaxy population than DESI, over a smaller area of sky.

Compared to ground-based observations, space observations afford higher angular resolution and a far lower NIR sky background. The *Euclid* and *WFIRST* (*Wide Field Infrared Survey Telescope*) missions will exploit these advantages, conducting large area imaging surveys for weak lensing and cluster studies and slitless spectroscopic surveys of emission-line galaxies for BAO and RSD studies. *WFIRST* also incorporates an imaging and spectrophotometric supernova (SN) survey, extending to redshift  $z \approx 1.7$ . Survey details are likely to evolve prior to launch, but in the current designs one can roughly characterize the difference between the *Euclid* and *WFIRST* dark energy experiments as “wide vs. deep,” with planned survey areas of 15,000 deg<sup>2</sup> and 2200 deg<sup>2</sup>, respectively. For weak lensing shape measurements, *Euclid* uses a single wide optical filter, while *WFIRST* uses three NIR filters. The *Euclid* galaxy redshift survey covers a large volume at relatively low space density, while the *WFIRST* survey provides denser sampling of structure in a smaller volume. There are numerous synergies among the LSST, *Euclid*, and *WFIRST* dark energy programs, as discussed in Ref. [26].

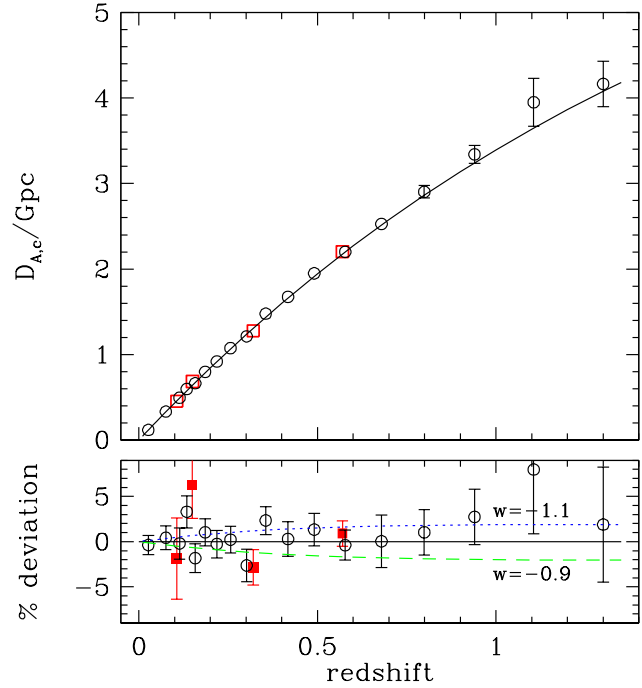
## 27.4. Current Constraints on Expansion, Growth, and Dark Energy

The last decade has seen dramatic progress in measurements of the cosmic expansion history and structure growth, leading to much tighter constraints on the parameters of dark energy models. CMB data from the *WMAP* and *Planck* satellites and from higher resolution ground-based experiments have provided an exquisitely detailed picture of structure at the recombination epoch and the first CMB-based measures of low redshift structure through lensing and SZ cluster counts. Cosmological supernova samples have increased in size from tens to many hundreds, with continuous coverage from  $z = 0$  to  $z \approx 1.4$ , alongside major improvements in data quality, analysis methods, and detailed understanding of local populations. BAO measurements have advanced from the first detections to 1–2% precision at multiple redshifts, with increasingly sophisticated methods for testing systematics, fitting models, and evaluating statistical errors. Constraints on low redshift structure from galaxy clusters have become more robust, with improved X-ray and SZ data and weak lensing mass calibrations, and they have been joined by the first precise structure constraints from cosmic shear weak lensing, galaxy-galaxy lensing, and redshift-space distortions. The precision of direct  $H_0$  measurements has sharpened from the roughly 10% error of the *HST* Key Project [27] to 3–4% in some recent analyses.

Our summary of current constraints here relies heavily on the analysis of Ref. [28], who combine BAO measurements, SN measurements, and *Planck* CMB data to examine a variety of dark energy models. While Ref. [28] uses the 2013 *Planck* data [29] rather than the 2015 data [30], we expect that changing to the 2015 data would make negligible difference to best-fit parameter values and only small changes to the statistical uncertainties on combined CMB+BAO+SN constraints. An analysis of dark energy and modified gravity models by the *Planck* team, using the 2015 *Planck* data and a somewhat different selection of low redshift data and model parameterizations, can be found in Ref. [31].

As an illustration of current measurements of the cosmic expansion history, Fig. 27.1 compares distance-redshift measurements from SN and BAO data to the predictions for a flat universe with a cosmological constant. SN cosmology relies on compilation analyses that try to bring data from different surveys probing distinct redshift ranges to a common scale. Here we use the “joint light curve analysis” (JLA) sample of Ref. [33], who carried out a careful intercalibration of the 3-year Supernova Legacy Survey (SNLS3, [34]) and the full SDSS-II Supernova Survey [35] data in combination with several local supernova samples and high-redshift supernovae from *HST*. Results from the Union2.1 sample [36], which partly overlaps JLA but has different analysis procedures, would be similar. For illustration purposes, we have binned the JLA data in redshift and plotted the diagonal elements of the covariance matrix as error bars, and we have converted the SN luminosity distances to an equivalent comoving angular diameter distance. Because the peak luminosity of a fiducial SN Ia is an unknown free parameter, the SN distance measurements could all be shifted up and down by a constant multiplicative factor; cosmological information resides in the relative distances as a function of redshift. For BAO data points we use the compilation of Ref. [28], taken from BAO analyses of the 6dFGS survey [37], the SDSS-II Main Galaxy Sample [38], and the LOWZ and CMASS galaxy samples of BOSS [39]. For the first three data points, values of  $D_V$  have been converted to  $D_{A,c}$ , while the CMASS data point uses the angular diameter distance measured directly from anisotropic BAO analysis. The BAO measurements are converted to absolute distances using the sound horizon scale  $r_s = 147.49$  Mpc from *Planck* 2013 CMB data, whose 0.4% uncertainty is small compared to the current BAO measurement errors.

The plotted cosmological model has  $\Omega_m = 0.308$  and  $h = 0.678$ , the best-fit values from *Planck* (TT+lowP+lensing) assuming  $w = -1$  and  $\Omega_{\text{tot}} = 1$  [32]. The SN, BAO, and CMB data sets, probing a wide range of redshifts with radically different techniques, are mutually consistent with the predictions of a flat  $\Lambda$ CDM cosmology. Other curves in the lower panel of Fig. 27.1 show the effect of changing  $w$  by  $\pm 0.1$  with all other parameters held fixed. However, such a

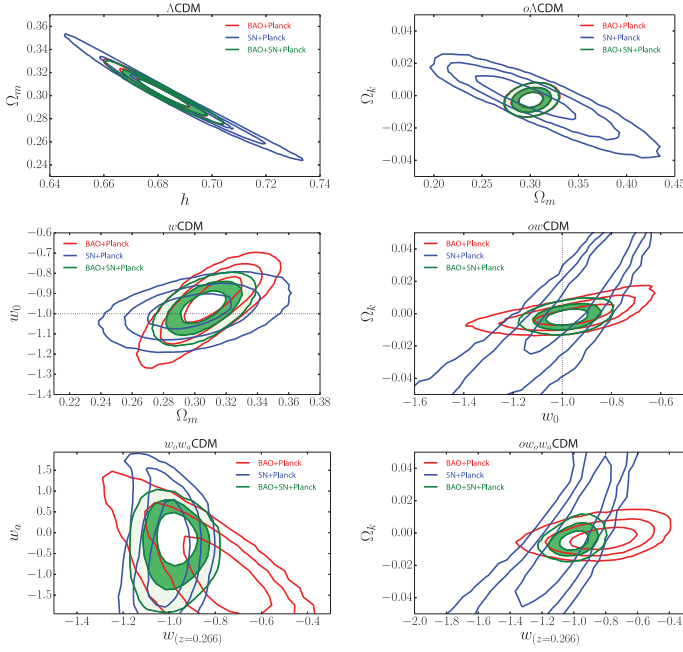


**Figure 27.1:** The distance-redshift relation measured from Type Ia SNe and BAO compared to the predictions (black curve) of a flat  $\Lambda$ CDM model with  $\Omega_m = 0.308$  and  $h = 0.678$ , the best-fit parameters inferred from *Planck* CMB data [32]. Circles show binned luminosity distances from the JLA SN sample [33], multiplied by  $(1+z)^{-1}$  to convert to comoving angular diameter distance. Squares show BAO distance measurements, converted to  $D_{A,c}(z)$  for the *Planck* cosmology and sound horizon, taken from Ref. [28]. The lower panel plots residuals from the  $\Lambda$ CDM prediction, with dashed and dotted curves that show the effect of changing  $w$  by  $\pm 0.1$  while all other parameters are held fixed. Note that the SN data points can be shifted up or down by a constant factor to account for freedom in the peak luminosity, while the BAO points are calibrated to 0.4% precision by the sound horizon scale computed from *Planck* data. In the upper panel, error bars are plotted only at  $z > 0.7$  to avoid visual confusion.

single-parameter comparison does not capture the impact of parameter degeneracies or the ability of complementary data sets to break them, and if one instead forced a match to CMB data by changing  $h$  and  $\Omega_m$  when changing  $w$  then the predicted BAO distances would diverge at  $z = 0$  rather than converging there.

Figure 27.2, taken directly from [28], shows cosmological parameter constraints in a series of models with increasingly flexible assumptions (from top left to bottom right) about dark energy and space curvature. These constraints use the BAO distance measurements shown in Fig. 27.1, with the separate  $D_{A,c}(z)$  and  $H(z)$  constraints from the BOSS CMASS sample at  $z = 0.57$ . They also include BAO constraints on  $D_{A,c}(z)$  and  $H(z)$  at  $z = 2.34$  from the BOSS Lyman- $\alpha$  forest as reported by Ref. [40]. They adopt the JLA SN data set plotted in Fig. 27.1, taking into account the full error covariance matrix reported by Ref. [33], which includes a detailed estimate of systematic uncertainties. The *Planck* CMB data are compressed into constraints on the baryon density  $\Omega_b h^2$ , the sum of baryon and CDM densities  $\Omega_m h^2$ , and the ratio  $D_{A,c}(1090)/r_s$  of the comoving angular diameter distance to redshift  $z = 1090$  divided by the sound horizon. Best-fit values and the  $3 \times 3$  covariance matrix of these quantities are determined from the public *Planck* likelihood chains. For the data combinations and models shown here, this compressed description captures the information content of the full CMB power spectrum almost perfectly; this would no longer be true when considering





**Figure 27.2:** Constraints on cosmological parameter combinations in a variety of dark energy models, taken from Ref. [28]. In each panel: red curves show 68%, 95%, and 99.7% confidence contours from BAO measurements with *Planck* CMB constraints; blue contours show the combination of SN measurements with *Planck* CMB; and green contours show the combination of all three, with the white zone interior to the dark green annulus marking the 68% confidence region. The upper left panel shows  $(\Omega_m, h)$  constraints assuming a flat universe with a cosmological constant. The upper right panel shows  $(\Omega_m, \Omega_K)$  constraints assuming a cosmological constant but allowing non-zero space curvature. The middle row shows constraints with the dark energy equation-of-state  $w_0$  as a (constant) free parameter, assuming a flat universe on the left and allowing non-zero curvature on the right. The bottom row shows the corresponding constraints for models with an evolving equation-of-state parameter  $w(a) = w_0 + w_a(1 - a)$ . In the bottom panels the  $x$ -axis quantity is the value of  $w$  at  $z = 0.266$ , the redshift at which it is best constrained by the full data combination in the flat universe model.

models with non-minimal neutrino mass or data sets that constrain the amplitude of matter clustering. Constraints from the full data combination on selected cosmological parameters for three dark energy models are listed in Table 27.3; this is a small subset of the models and data combinations reported in table IV of Ref. [28].

There are numerous points to take away from Fig. 27.2 and Table 27.3. For the flat  $\Lambda$ CDM model, the combination of CMB and BAO data provides tight constraints on parameters, as discussed at greater length in Sec. 25 of this *Review*. Assuming a cosmological constant, the CMB+BAO combination yields a tight constraint on space curvature consistent with a flat universe, implying  $\Omega_{\text{tot}} = 1.002 \pm 0.003$ . The addition of SN data does not tighten the constraints in cosmological constant models (top row), but it significantly tightens constraints in models that allow an evolving dark energy density. In all of the more flexible models, the parameter values of flat  $\Lambda$ CDM ( $w = -1$ ,  $w_a = 0$ ,  $\Omega_K = 0$ ) lie within the 68% confidence region of the full CMB+BAO+SN combination. Even with the flexibility of an evolving equation of state governed by Eq. (27.4), curvature is tightly constrained by the full data combination. For a constant equation of state, the error on  $w$  is  $\approx 0.05$ , and even in the  $w_0 - w_a$  model the value of  $w$  at the pivot redshift  $z_p = 0.266$  is constrained to  $\pm 0.05$ . However, the full CMB+BAO+SN data combination still provides only weak constraints on evolution of the

**Table 27.3:** Constraints on parameters (68% confidence limits) from the combination of BAO, SN, and CMB data as reported by Ref. [28], for three choices of model assumptions: constant  $w$  with a flat universe, constant  $w$  with free space curvature, and evolving  $w$  with a flat universe. In the third model, the constraint on  $w$  is reported at  $z = 0.266$ , where it is best constrained.

Parameter	Model		
	$w$ CDM (flat)	$ow$ CDM	$w_0w_a$ CDM (flat)
$w$	$-0.97 \pm 0.05$	$-0.98 \pm 0.06$	$-0.97 \pm 0.05$
$w_a$	0 (assumed)	0 (assumed)	$-0.2 \pm 0.4$
$\Omega_m$	$0.305 \pm 0.010$	$0.303 \pm 0.010$	$0.307 \pm 0.011$
$\Omega_{\text{tot}}$	1.0 (assumed)	$1.002 \pm 0.003$	1.0 (assumed)
$h$	$0.676 \pm 0.011$	$0.676 \pm 0.011$	$0.676 \pm 0.011$
$\sigma_8(\Omega_m/0.30)^{0.4}$	$0.811 \pm 0.021$	$0.805 \pm 0.022$	$0.821 \pm 0.030$

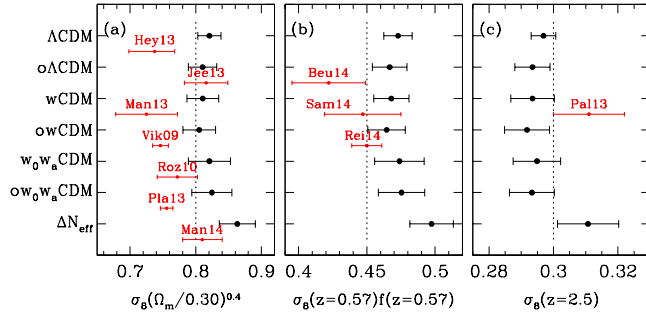
equation of state, allowing  $w_a = -0.2 \pm 0.4$  even when assuming a flat universe.

As discussed by [28], the flat  $\Lambda$ CDM model provides a statistically good fit to the CMB+BAO+SN data combination presented here. However, the Lyman- $\alpha$  forest BAO measurements at  $z \approx 2.3$  disagree with the model predictions at the  $\approx 2.5\sigma$  level [40]. None of the more flexible models illustrated in Fig. 27.2 significantly reduces this tension, and Ref. [28] considers a variety of more elaborate models (decaying dark matter, early dark energy, massive neutrinos, additional relativistic species) that also fail to remove it. The lack of a plausible alternative model, and the acceptable total  $\chi^2$  when all data points are considered equally, suggests that the discrepancy with Lyman- $\alpha$  forest BAO is either a statistical fluke or an unrecognized systematic bias in the measurement. This remains an interesting area for future investigation, as a tightening of error bars without a change in central value would imply a breakdown of this entire class of dark energy models at  $z \approx 2 - 3$ , or an unanticipated astrophysical effect on the imprint of BAO in the Lyman- $\alpha$  forest.

The 2014 edition of this review highlighted two areas of tension between predictions of the flat  $\Lambda$ CDM model and low-redshift observations: distance-ladder measurements of  $H_0$  and weak lensing or cluster estimates of matter fluctuations. A  $\Lambda$ CDM fit to *Planck* data alone predicts  $H_0 = 67.8 \pm 0.9 \text{ km s}^{-1} \text{ Mpc}^{-1}$  (see Chapter 28 of this *Review*). This is lower than most recent determinations of  $H_0$  that use *HST* observations of Cepheid variables in external galaxies to calibrate secondary distance indicators, particularly Type Ia SNe, which can in turn measure distances to galaxies in the Hubble flow. For example, Ref. [41] finds  $H_0 = 73.8 \pm 2.4 \text{ km s}^{-1} \text{ Mpc}^{-1}$  and Ref. [42] finds  $H_0 = 74.3 \pm 2.1 \text{ km s}^{-1} \text{ Mpc}^{-1}$ , with both groups including an estimate of systematic uncertainties in their error budgets. However, Ref. [43], reanalyzing the data set of Ref. [41] with a different treatment of outliers, argues for a lower central value and larger error bars, which together reduce the tension with *Planck*+ $\Lambda$ CDM below  $2\sigma$  significance. More recently, Ref. [44] have argued that correcting the Ref. [41] value for an offset of SNIa luminosities between star-forming and passive environments lowers the inferred  $H_0$  to  $70.6 \pm 2.6 \text{ km s}^{-1} \text{ Mpc}^{-1}$ , consistent with the CMB at the  $1\sigma$  level.

Another recent development is the “inverse distance ladder” determination of  $H_0$  by Ref. [28], who combine the BAO and SN data shown in Fig. 27.1 with the *Planck*-calibrated value of the sound horizon scale,  $r_s = 147.49 \pm 0.59 \text{ Mpc}$ . The CMB-only prediction of  $H_0$  depends critically on the assumptions of a flat universe and a cosmological constant, and loosening either assumption allows a much wider range of  $H_0$ . The method of Ref. [28], by contrast, is insensitive to assumptions about flatness or dark energy, because BAO provide precise absolute distance measurements at  $z = 0.3 - 0.6$ , and the high-precision relative distance scale from SNe transfers this absolute measurement to  $z = 0$ , using empirical data instead of an adopted cosmological model. Even allowing a very flexible dark energy parameterization and non-zero space curvature, Ref. [28] obtains 1.7% precision on  $H_0$ , with a value

$H_0 = 67.3 \pm 1.1 \text{ km s}^{-1} \text{ Mpc}^{-1}$  in essentially perfect agreement with the *Planck*+ $\Lambda$ CDM prediction. These measurements could still be reconciled with  $H_0 \geq 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$  by altering the *pre-recombination* physics of the standard model in a way that shrinks the BAO standard ruler, for instance by adding extra relativistic degrees of freedom. However, it seems increasingly unlikely that the Cepheid-based measurements of  $H_0$  are telling us something surprising about the late time behavior of dark energy, and more likely that they simply overestimate the true value.



**Figure 27.3:** Comparison of observational estimates of matter clustering (red points) to the amplitude predicted for a variety of dark energy models constrained by CMB+BAO+SN data (black points) at  $z \approx 0$  (left),  $z = 0.57$  (middle), and  $z = 2.5$  (right), taken from Ref. [28]. Black points with error bars correspond to the 68% confidence range of predictions for the model indicated on the left axis. (Models beginning “ $\phi$ ” allow non-zero curvature, while other models assume a flat universe.) Fractional errors for the red points are taken from the observational references given in Ref. [28], and the vertical placement of these points is arbitrary. The observational estimates of  $\sigma_8\Omega_m^{0.4}$  in the left panel come from a variety of weak lensing and cluster studies; the estimates of  $\sigma_8(z)f(z)$  in the middle panel come from RSD analyses of the BOSS CMASS galaxy sample; and the estimate of  $\sigma_8(z = 2.5)$  comes from the 1-dimensional power spectrum of the BOSS Lyman- $\alpha$  forest.

The amplitude of CMB anisotropies is proportional to the amplitude of density fluctuations present at recombination, and by assuming GR and a specified dark energy model one can extrapolate the growth of structure forward to the present day to predict  $\sigma_8$ . As discussed in Sec. 27.3, probes of low redshift structure typically constrain the combination  $\sigma_8\Omega_m^\alpha$  with  $\alpha \approx 0.3$ –0.5. Figure 27.3, taken from Ref. [28], compares predictions of low redshift clustering (black points) from models constrained by CMB+BAO+SN to a variety of observational estimates (red points). The model assumed for each prediction is indicated on the left axis. In the left panel, estimates of  $\sigma_8(\Omega_m/0.30)^{0.4}$  at  $z \approx 0$  come from cosmic shear (points labeled Hey13, Jee13 in Fig. 27.3), from galaxy-galaxy lensing (Man13), and from clusters (Vik09, Roz10, Pla13, Man14; see Ref. [28] for the observational references). In the middle panel, the values of  $\sigma_8(z = 0.57)f(z = 0.57)$  come from three RSD analyses of the BOSS galaxy survey; these analyses use different modeling methods but examine largely the same data. In the right panel the estimate of  $\sigma_8(z = 2.5)$  comes from modeling the 1-dimensional power spectrum of the BOSS Lyman- $\alpha$  forest. In the left panel, many but not all of the estimates lie below the model predictions. A straight unweighted average of the observational data points yields  $\sigma_8(\Omega_m/0.3)^{0.4} = 0.766 \pm 0.012$ , while the flat  $\Lambda$ CDM prediction is  $0.821 \pm 0.018$ . This difference is  $\approx 2\sigma$ , but the key question is whether some of the estimates are systematically biased, and if so which ones. In the middle panel, the RSD growth estimates again lie below the model predictions, but the observational uncertainties are too large to draw an interesting conclusion. At  $z = 2.5$ , on the other hand, the fluctuation amplitude inferred from the Lyman- $\alpha$  forest is (slightly) above the model prediction.

Relative to our 2014 *Review* (which compared  $\Lambda$ CDM to the

constraints labeled here as Hey13, Vik09, Roz10, and Pla13), the addition of new data has made the case for a conflict in matter clustering weaker, or at least more confused. The 2015 *Planck* data add two further ingredients to this discussion. First, they confirm the high normalization of  $(\sigma_8, \Omega_m)$  relative to earlier values from *WMAP*, indicating that the high model predictions are not a statistical fluctuation in early *Planck* data or a systematic error in the 2013 analysis. (The 2015 *Planck* analysis does change in some significant ways, but the net impact on  $\sigma_8$  and  $\Omega_m$  is small.) Second, CMB lensing in *Planck* 2015 yields a roughly 3% measurement of the matter clustering amplitude over an effective redshift range  $z \approx 2$ –5, and this measurement is in excellent agreement with the flat  $\Lambda$ CDM prediction. The CMB lensing and Lyman- $\alpha$  forest measurements imply that deviation from GR-predicted structure growth, if it occurs, must set in mainly at  $z < 2$ . A low redshift onset would not necessarily be surprising, however, as it would coincide with the era of cosmic acceleration.

## 27.5. Summary and Outlook

Figure 27.2 and Table 27.3 focus on model parameter constraints, but as a description of the observational situation it is most useful to characterize the precision, redshift range, and systematic uncertainties of the basic expansion and growth measurements. At present, supernova surveys constrain distance ratios at the 1–2% level in redshift bins of width  $\Delta z = 0.1$  over the range  $0 < z < 0.6$ , with larger but still interesting error bars out to  $z \approx 1.2$ . These measurements are currently limited by systematics tied to photometric calibration, extinction, and reddening, host galaxy correlations, and possible evolution of the SN population. BAO surveys have measured the absolute distance scale (calibrated to the sound horizon  $r_s$ ) to 4% at  $z = 0.15$ , 2% at  $z = 0.32$ , 1% at  $z = 0.57$ , and 2% at  $z = 2.3$ . Multiple studies have used clusters of galaxies or weak lensing cosmic shear or galaxy-galaxy lensing to measure a parameter combination  $\sigma_8\Omega_m^\alpha$  with  $\alpha \approx 0.3$ –0.5. The estimated errors of these studies, including both statistical contributions and identified systematic uncertainties, are about 5%. RSD measurements constrain the combination  $f(z)\sigma_8(z)$ , with recent determinations spanning the redshift range  $0 < z < 0.9$  with typical estimated errors of about 10%. These errors are dominated by statistics, but shrinking them further will require improvements in modeling non-linear effects on small scales. Direct distance-ladder estimates of  $H_0$  now span a small range (using overlapping data but distinct treatments of key steps), with individual studies quoting uncertainties of 3–5%, with similar statistical and systematic contributions. *Planck* data and higher resolution ground-based experiments now measure CMB anisotropy with exquisite precision; for example, CMB measurements now constrain the physical size of the BAO sound horizon to 0.3% and the angular scale of the sound horizon to 0.01%.

A flat  $\Lambda$ CDM model with standard radiation and neutrino content can fit the CMB data and the BAO and SN distance measurements to within their estimated uncertainties, excepting a moderately significant discrepancy for Lyman- $\alpha$  forest BAO at  $z = 2.3$ . However the CMB+BAO parameters for this model are in approximately  $2\sigma$  tension with some of the direct  $H_0$  measurements and many but not all of the cluster and weak lensing analyses, disagreeing by about 10% in each case. Agreement of the “inverse distance ladder” value of  $H_0$  with the *Planck*+ $\Lambda$ CDM value suggests that the current direct measurements are systematically high. Alternatively, a change to pre-recombination physics (such as extra relativistic energy density) could shrink the BAO standard ruler and raise the inferred  $H_0$ , but changes large enough to allow  $H_0 \geq 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$  might run afoul of the CMB power spectrum shape. CMB lensing and Lyman- $\alpha$  forest measurements show good agreement with  $\Lambda$ CDM-predicted structure growth at  $z \approx 2$ –4, so if the discrepancies with lower redshift measurements are real then the deviations in growth must set in at late times. At present, none of the tensions in the data provide compelling evidence for new physics. Moving forward, the community will have to balance the requirement of strong evidence for interesting claims (such as  $w \neq -1$  or deviations from GR) against the danger of confirmation bias, *i.e.*, discounting observations or error estimates when they do not overlap simple theoretical expectations.



There are many ongoing projects that should lead to improvement in observational constraints in the near-term and over the next 15 years, as summarized above in Table 27.2. Final analyses of *Planck* temperature, polarization, and CMB lensing maps will improve estimates of the electron scattering optical depth and tighten other parameter constraints, thus sharpening tests based on structure growth. Preliminary results suggest a small reduction in the inferred  $\sigma_8$ , which goes in the direction of reducing tensions. Final analyses of BOSS will slightly reduce BAO errors at  $z < 0.6$  and shed light on the significance of the Lyman- $\alpha$  forest tension at  $z = 2.3$ . Its successor eBOSS will yield the first BAO measurements in the redshift range  $1 < z < 2$  and improved precision at lower and higher redshifts. The HETDEX project will measure BAO with Lyman- $\alpha$  emission line galaxies at  $z = 2-3$ , providing an independent check on Lyman- $\alpha$  forest results with completely different structure tracers. The same galaxy surveys carried out for BAO also provide data for RSD measurements of structure growth and AP measurements of cosmic geometry. With improved theoretical modeling there is potential for substantial precision gains over current constraints from these methods. DES, which started operations in August 2013 and will run through 2018, will provide a sample of several thousand Type Ia SNe, enabling smaller statistical errors and division of the sample into subsets for cross-checking evolutionary effects and other systematics. DES imaging will be similar in depth but 50 times larger in area than CFHTLenS, providing a much more powerful weak lensing data set and weak lensing mass calibration of enormous samples of galaxy clusters (tens of thousands). Weak lensing surveys from HSC on the Subaru telescope will be smaller in area but deeper, with a comparable number of lensed galaxies. These new weak lensing data sets hold the promise of providing structure growth constraints at the same (roughly 1%) level of precision as the best current expansion history constraints, allowing a much more comprehensive test of cosmic acceleration models. Controlling measurement and modeling systematics at the level demanded by these surveys' statistical power will be a major challenge, but the payoff in improved precision is large. Uncertainties in direct determinations of  $H_0$  should be reduced by further observations with *HST* and, in the longer run, by Cepheid parallaxes from the *GAIA* mission, by the ability of the *James Webb Space Telescope* to discover Cepheids in more distant SN Ia calibrator galaxies, and by independent estimates from larger samples of maser galaxies and gravitational lensing time delays.

A still more ambitious period begins late in this decade and continues through the 2020s, with experiments that include DESI, Subaru PFS, LSST, and the space missions *Euclid* and *WFIRST*. DESI and PFS both aim for major improvements in the precision of BAO, RSD, and other measurements of galaxy clustering in the redshift range  $0.8 < z < 2$ , where large comoving volume allows much smaller cosmic variance errors than low redshift surveys like BOSS. LSST will be the ultimate ground-based optical weak lensing experiment, measuring several billion galaxy shapes over 20,000 deg<sup>2</sup> of the southern hemisphere sky, and it will detect and monitor many thousands of SNe per year. *Euclid* and *WFIRST* also have weak lensing as a primary science goal, taking advantage of the high angular resolution and extremely stable image quality achievable from space. Both missions plan large spectroscopic galaxy surveys, which will provide better sampling at high redshifts than DESI or PFS because of the lower infrared sky background above the atmosphere. *WFIRST* is also designed to carry out what should be the ultimate supernova cosmology experiment, with deep, high resolution, near-IR observations and the stable calibration achievable with a space platform.

Performance forecasts necessarily become more uncertain the further ahead we look, but collectively these experiments are likely to achieve 1–2 order of magnitude improvements over the precision of current expansion and growth measurements, while simultaneously extending their redshift range, improving control of systematics, and enabling much tighter cross-checks of results from entirely independent methods. The critical clue to the origin of cosmic acceleration could also come from a surprising direction, such as laboratory or solar system tests that challenge GR, time variation of fundamental “constants,” or anomalous behavior of gravity in some astronomical

environments. Experimental advances along these multiple axes could confirm today's relatively simple, but frustratingly incomplete, “standard model” of cosmology, or they could force yet another radical revision in our understanding of energy, or gravity, or the spacetime structure of the Universe.

## References:

1. A. Einstein, Sitzungsber. Preuss. Akad. Wiss. Berlin (Math. Phys.), 142 (1917).
2. Y.B. Zeldovich, Soviet Physics Uspekhi **11**, 381 (1968).
3. A. Friedmann, On the curvature of space. Z. Phys. **10**, 377 (1922).
4. G. Lemaître, Annales de la Societe Scietifique de Bruxelles **47**, 49 (1927).
5. E. Hubble, Proc. Nat. Acad. Sci. **15**, 168 (1929).
6. A. Einstein and W. de Sitter, Proc. Nat. Acad. Sci. **18**, 213 (1932).
7. For background and definitions, see Big-Bang Cosmology – Sec. 22 of this *Review*.
8. A.G. Riess *et al.* [Supernova Search Team Collab.], Astron. J. **116**, 1009 (1998).
9. S. Perlmutter *et al.*, [Supernova Cosmology Project Collab.], Astrophys. J. **517**, 565 (1999).
10. P. de Bernardis *et al.* [Boomerang Collab.], Nature **404**, 955 (2000).
11. S. Hanany *et al.*, Astrophys. J. **545**, L5 (2000).
12. D.H. Weinberg *et al.*, Phys. Reports **530**, 87 (2013).
13. B. Ratra and P.J.E. Peebles, Phys. Rev. **D37**, 3406 (1988) C. Wetterich, Nucl. Phys. **B302**, 668 (1988).
14. Excellent overviews of the theory and phenomenology of modified gravity models can be found in the review articles of B. Jain and J. Khoury, Ann. Phys. **325**, 1479 (2010) and A. Joyce *et al.*, Phys. Reports **568**, 1 (2015).
15. S.M. Carroll *et al.*, Phys. Rev. **D70**, 043528 (2004).
16. G.R. Dvali, G. Gabadadze, and M. Porrati, Phys. Lett. **B485**, 208 (2000).
17. C.M. Will, Living Reviews in Relativity, **9**, 3 (2006). See also the chapter on Experimental Tests of Gravitational Theory — in this *Review*.
18. B. Jain, V. Vikram, and J. Sakstein, Astrophys. J. **779**, 39 (2013) J. Wang, L. Hui, and J. Khoury, Phys. Rev. Lett. **109**, 241301 (2012).
19. Multiple investigations including M. Fairbairn and A. Goobar, Phys. Lett. **B642**, 432 (2006); Y.-S. Song, I. Sawicki, and W. Hu, Phys. Rev. **D75**, 064003 (2007); C. Blake *et al.*, Mon. Not. Roy. Astron. Soc. **415**, 2876 (2011).
20. E.V. Linder, Phys. Rev. **D72**, 043529 (2005).
21. This is essentially the FoM proposed in the Dark Energy Task Force (DETF) report, A. Albrecht *et al.*, astro-ph/0609591, though they based their FoM on the area of the 95% confidence ellipse in the  $w_0 - w_a$  plane.
22. For high accuracy, the impact of acoustic oscillations must be computed with a full Boltzmann code, but the simple integral for  $r_s$  captures the essential physics and the scaling with cosmological parameters.
23. R.A. Sunyaev and Y.B. Zeldovich, Astrophys. Space Sci. **7**, 3 (1970).
24. C. Alcock and B. Paczynski, Nature **281**, 358 (1979).
25. D. Weinberg *et al.*, Snowmass 2013 report on Facilities for Dark Energy Investigations, arXiv:1309.5380.
26. B. Jain *et al.*, The Whole is Greater than the Sum of the Parts: Optimizing the Joint Science Return from LSST, Euclid and WFIRST, arXiv:1501.07897.
27. W.L. Freedman *et al.*, Astrophys. J. **553**, 47 (2001).
28. E. Aubourg *et al.*, Phys. Rev. **D92**, 123516 (2015).
29. Planck Collab. 2013 Results XVI, Astron. & Astrophys. **571**, A16 (2014).
30. Planck Collab. 2015 Results I, arXiv:1502.01582.
31. Planck Collab. 2015 Results XIV, arXiv:1502.01590.
32. Planck Collab. 2015 Results XIII, arXiv:1502.01589.
33. M. Betoule *et al.*, Astron. & Astrophys. **568**, 22 (2014).
34. M. Sullivan *et al.*, Astrophys. J. **737**, 102 (2011).

- 35. J.A. Frieman *et al.*, *Astrophys. J.* **135**, 338 (2008).
- 36. N. Suzuki *et al.*, *Astrophys. J.* **746**, 85 (2012).
- 37. F. Beutler *et al.*, *Mon. Not. Roy. Astron. Soc.* **416**, 3017 (2011).
- 38. A.J. Ross *et al.*, *Mon. Not. Roy. Astron. Soc.* **449**, 835 (2015).
- 39. L. Anderson *et al.*, *Mon. Not. Roy. Astron. Soc.* **441**, 24 (2014).
- 40. T. Delubac *et al.*, *Astron. & Astrophys.* **574**, 59 (2015).
- 41. A.G. Riess *et al.*, *Astrophys. J.* **730**, 119 (2011).
- 42. W.L. Freedman *et al.*, *Astrophys. J.* **758**, 24 (2012).
- 43. G. Efstathiou, *Mon. Not. Roy. Astron. Soc.* **440**, 1138 (2014).
- 44. M. Rigault *et al.*, *Astrophys. J.* **802**, 20 (2015).

## 28. COSMIC MICROWAVE BACKGROUND

Revised September 2015 by D. Scott (University of British Columbia) and G.F. Smoot (UCB/LBNL).

### 28.1. Introduction

The energy content in radiation from beyond our Galaxy is dominated by the cosmic microwave background (CMB), discovered in 1965 [1]. The spectrum of the CMB is well described by a blackbody function with  $T = 2.7255$  K. This spectral form is a main supporting pillar of the hot Big Bang model for the Universe. The lack of any observed deviations from a blackbody spectrum constrains physical processes over cosmic history at redshifts  $z \lesssim 10^7$  (see earlier versions of this review).

Currently the key CMB observable is the angular variation in temperature (or intensity) correlations, and now to some extent polarization [2]. Since the first detection of these anisotropies by the Cosmic Background Explorer (*COBE*) satellite [3], there has been intense activity to map the sky at increasing levels of sensitivity and angular resolution by ground-based and balloon-borne measurements. These were joined in 2003 by the first results from NASA's Wilkinson Microwave Anisotropy Probe (*WMAP*) [4], which were improved upon by analyses of the 3-year, 5-year, 7-year, and 9-year *WMAP* data [5,6,7,8]. In 2013 we had the first results [9] from the third generation CMB satellite, ESA's *Planck* mission [10,11], now enhanced by results from the the 2015 *Planck* data release [12,13]. Additionally, CMB anisotropies have been extended to smaller angular scales by ground-based experiments, particularly the Atacama Cosmology Telescope (ACT) [14] and the South Pole Telescope (SPT) [15]. Together these observations have led to a stunning confirmation of the 'Standard Model of Cosmology.' In combination with other astrophysical data, the CMB anisotropy measurements place quite precise constraints on a number of cosmological parameters, and have launched us into an era of precision cosmology. As the CMB turns 50 and the program to map temperature anisotropies is wrapping up, attention is increasingly focussing on polarization measurements as the future arena in which to test fundamental physics.

### 28.2. CMB Spectrum

It is well known that the spectrum of the microwave background is very precisely that of blackbody radiation, whose temperature evolves with redshift as  $T(z) = T_0(1+z)$  in an expanding universe. As a direct test of its cosmological origin, this relationship is being tested by measuring the strengths of emission and absorption lines in high-redshift systems [16].

Measurements of the spectrum are consistent with a blackbody distribution over more than three decades in frequency (there is a claim by ARCADE [17] of a possible unexpected extragalactic emission signal at low frequency, but the interpretation is debated [18]). All viable cosmological models predict a very nearly Planckian spectrum to within the current observational limits. Because of this, measurements of deviations from a blackbody spectrum have received little attention in recent years, with only a few exceptions. However, that situation may be about to change, since proposed experiments (such as PIXIE [19]) have the potential to dramatically improve the constraints on energy release in the early Universe. It now seems feasible to probe spectral distortion mechanisms that are *required* in the standard picture, such as those arising from the damping and dissipation of relatively small primordial perturbations, or the average effect of inverse Compton scattering. A more ambitious goal would be to reach the precision needed to detect the residual lines from the cosmological recombination of hydrogen and helium and hence test whether conditions at  $z \gtrsim 1000$  accurately follow those in the standard picture [20].

### 28.3. Description of CMB Anisotropies

Observations show that the CMB contains temperature anisotropies at the  $10^{-5}$  level and polarization anisotropies at the  $10^{-6}$  (and lower) level, over a wide range of angular scales. These anisotropies are usually expressed by using a spherical harmonic expansion of the CMB sky:

$$T(\theta, \phi) = \sum_{\ell m} a_{\ell m} Y_{\ell m}(\theta, \phi)$$

(with the linear polarization pattern written in a similar way using the so-called spin-2 spherical harmonics). Increasing angular resolution requires that the expansion goes to higher and higher multipoles. Because there are only very weak phase correlations seen in the CMB sky and since we notice no preferred direction, the vast majority of the cosmological information is contained in the temperature 2-point function, *i.e.*, the variance as a function of angular separation. Equivalently, the power per unit  $\ln \ell$  is  $\ell \sum_m |a_{\ell m}|^2 / 4\pi$ .

#### 28.3.1. The Monopole :

The CMB has a mean temperature of  $T_\gamma = 2.7255 \pm 0.0006$  K ( $1\sigma$ ) [21], which can be considered as the monopole component of CMB maps,  $a_{00}$ . Since all mapping experiments involve difference measurements, they are insensitive to this average level; monopole measurements can only be made with absolute temperature devices, such as the FIRAS instrument on the *COBE* satellite [22]. The measured  $kT_\gamma$  is equivalent to  $0.234$  meV or  $4.60 \times 10^{-10} m_e c^2$ . A blackbody of the measured temperature has a number density  $n_\gamma = (2\zeta(3)/\pi^2) T_\gamma^3 \simeq 411 \text{ cm}^{-3}$ , energy density  $\rho_\gamma = (\pi^2/15) T_\gamma^4 \simeq 4.64 \times 10^{-34} \text{ g cm}^{-3} \simeq 0.260 \text{ eV cm}^{-3}$ , and a fraction of the critical density  $\Omega_\gamma \simeq 5.38 \times 10^{-5}$ .

#### 28.3.2. The Dipole :

The largest anisotropy is in the  $\ell = 1$  (dipole) first spherical harmonic, with amplitude  $3.3645 \pm 0.0020$  mK [12]. The dipole is interpreted to be the result of the Doppler boosting of the monopole caused by the solar system motion relative to the nearly isotropic blackbody field, as broadly confirmed by measurements of the radial velocities of local galaxies (*e.g.*, Ref. [23]). The motion of an observer with velocity  $\beta \equiv v/c$  relative to an isotropic Planckian radiation field of temperature  $T_0$  produces a Lorentz-boosted temperature pattern

$$\begin{aligned} T(\theta) &= T_0(1 - \beta^2)^{1/2} / (1 - \beta \cos \theta) \\ &\simeq T_0 \left[ 1 + \beta \cos \theta + \left( \beta^2/2 \right) \cos 2\theta + \mathcal{O}(\beta^3) \right]. \end{aligned}$$

At every point in the sky, one observes a blackbody spectrum, with temperature  $T(\theta)$ . The spectrum of the dipole has been confirmed to be the differential of a blackbody spectrum [24]. At higher order there are additional effects arising from aberration and from modulation of the anisotropy pattern, which have also been observed [25].

The implied velocity for the solar system barycenter is  $v = 370.09 \pm 0.22 \text{ km s}^{-1}$ , assuming a value  $T_0 = T_\gamma$ , towards  $(l, b) = (264.00^\circ \pm 0.03^\circ, 48.24^\circ \pm 0.02^\circ)$  [12]. Such a solar system motion implies a velocity for the Galaxy and the Local Group of galaxies relative to the CMB. The derived value is  $v_{\text{LG}} = 627 \pm 22 \text{ km s}^{-1}$  towards  $(l, b) = (276^\circ \pm 3^\circ, 30^\circ \pm 3^\circ)$  [26], where most of the error comes from uncertainty in the velocity of the solar system relative to the Local Group.

The dipole is a frame-dependent quantity, and one can thus determine the 'absolute rest frame' as that in which the CMB dipole would be zero. Our velocity relative to the Local Group, as well as the velocity of the Earth around the Sun, and any velocity of the receiver relative to the Earth, is normally removed for the purposes of CMB anisotropy study. The dipole is now routinely used as a primary calibrator for mapping experiments, either via the time-varying orbital motion of the Earth, or through the cosmological dipole measured by satellite experiments.

#### 28.3.3. Higher-Order Multipoles :

The variations in the CMB temperature maps at higher multipoles ( $\ell \geq 2$ ) are interpreted as being mostly the result of perturbations in the density of the early Universe, manifesting themselves at the epoch of the last scattering of the CMB photons. In the hot Big Bang picture, the expansion of the Universe cools the plasma so that by a redshift  $z \simeq 1100$  (with little dependence on the details of the model), the hydrogen and helium nuclei can bind electrons into neutral atoms, a process usually referred to as recombination [27]. Before this epoch, the CMB photons were tightly coupled to the baryons, while afterwards they could freely stream towards us. By measuring the  $a_{\ell m}$ s we are thus learning directly about physical conditions in the early Universe.

A statistically isotropic sky means that all *ms* are equivalent, *i.e.*, there is no preferred axis, so that the temperature correlation function

between two positions on the sky depends only on angular separation and not orientation. Together with the assumption of Gaussian statistics (*i.e.*, no correlations between the modes), the variance of the temperature field (or equivalently the power spectrum in  $\ell$ ) then fully characterizes the anisotropies. The power summed over all  $m$ s at each  $\ell$  is  $(2\ell+1)C_\ell/(4\pi)$ , where  $C_\ell \equiv \langle |a_{\ell m}|^2 \rangle$ . Thus averages of  $a_{\ell m}$ s over  $m$  can be used as estimators of the  $C_\ell$ s to constrain their expectation values, which are the quantities predicted by a theoretical model. For an idealized full-sky observation, the variance of each measured  $C_\ell$  (*i.e.*, the variance of the variance) is  $[2/(2\ell+1)]C_\ell^2$ . This sampling uncertainty (known as ‘cosmic variance’) comes about because each  $C_\ell$  is  $\chi^2$  distributed with  $(2\ell+1)$  degrees of freedom for our observable volume of the Universe. For fractional sky coverage,  $f_{\text{sky}}$ , this variance is increased by  $1/f_{\text{sky}}$  and the modes become partially correlated.

It is important to understand that theories predict the expectation value of the power spectrum, whereas our sky is a single realization. Hence the cosmic variance is an unavoidable source of uncertainty when constraining models; it dominates the scatter at lower  $\ell$ s, while the effects of instrumental noise and resolution dominate at higher  $\ell$ s [28].

Theoretical models generally predict that the  $a_{\ell m}$  modes are Gaussian random fields to high precision, matching the empirical tests, *e.g.*, standard slow-roll inflation’s non-Gaussian contribution is expected to be at least an order of magnitude below current observational limits [29]. Although non-Gaussianity of various forms is possible in early Universe models, tests show that Gaussianity is an extremely good simplifying approximation [30]. The only current indications of any non-Gaussianity or statistical anisotropy are some relatively weak signatures at large scales, seen in both *WMAP* [31] and *Planck* data [32], but not of high enough significance to reject the simplifying assumption. Nevertheless, models that deviate from the inflationary slow-roll conditions can have measurable non-Gaussian signatures. So while the current observational limits make the power spectrum the dominant probe of cosmology, it is worth noting that higher-order correlations are beginning to be a tool for constraining otherwise viable theories.

#### 28.3.4. Angular Resolution and Binning :

There is no one-to-one conversion between multipole  $\ell$  and the angle subtended by a particular spatial scale projected onto the sky. However, a single spherical harmonic  $Y_{\ell m}$  corresponds to angular variations of  $\theta \sim \pi/\ell$ . CMB maps contain anisotropy information from the size of the map (or in practice some fraction of that size) down to the beam-size of the instrument,  $\sigma$  (the standard deviation of the beam, in radians). One can think of the effect of a Gaussian beam as rolling off the power spectrum with the function  $e^{-\ell(\ell+1)\sigma^2}$ .

For less than full sky coverage, the  $\ell$  modes become correlated. Hence, experimental results are usually quoted as a series of ‘band powers,’ defined as estimators of  $\ell(\ell+1)C_\ell/2\pi$  over different ranges of  $\ell$ . Because of the strong foreground signals in the Galactic Plane, even ‘all-sky’ surveys, such as *WMAP* and *Planck* involve a cut sky. The amount of binning required to obtain uncorrelated estimates of power also depends on the map size.

### 28.4. Cosmological Parameters

The current ‘Standard Model’ of cosmology contains around 10 free parameters (see The Cosmological Parameters—Sec. 25 of this *Review*). The basic framework is the Friedmann-Robertson-Walker (FRW) metric (*i.e.*, a universe that is approximately homogeneous and isotropic on large scales), with density perturbations laid down at early times and evolving into today’s structures (see Big-Bang cosmology—Sec. 22 of this *Review*). The most general possible set of density variations is a linear combination of an adiabatic density perturbation and some isocurvature perturbations. Adiabatic means that there is no change to the entropy per particle for each species, *i.e.*,  $\delta\rho/\rho$  for matter is  $(3/4)\delta\rho/\rho$  for radiation. Isocurvature means that the set of individual density perturbations adds to zero, for example, matter perturbations compensate radiation perturbations so that the total energy density remains unperturbed, *i.e.*,  $\delta\rho$  for matter is  $-\delta\rho$  for radiation. These different modes give rise to distinct (temporal) phases during growth, with those of the adiabatic scenario looking exactly like the data. Models that generate mainly isocurvature type

perturbations (such as most topological defect scenarios) are no longer considered to be viable. However, an admixture of the adiabatic mode with up to about 4% isocurvature contribution (depending on details of the mode) is still allowed [33,34].

#### 28.4.1. Initial Condition Parameters :

Within the adiabatic family of models, there is, in principle, a free function describing the variation of comoving curvature perturbations,  $\mathcal{R}(\mathbf{x}, t)$ . The great virtue of  $\mathcal{R}$  is that it is constant in time for a purely adiabatic perturbation. There are physical reasons to anticipate that the variance of these perturbations will be described well by a power law in scale, *i.e.*, in Fourier space  $\langle |\mathcal{R}|_k^2 \rangle \propto k^{n_s-4}$ , where  $k$  is wavenumber and  $n_s$  is the usual definition of spectral index. So-called ‘scale-invariant’ initial conditions (meaning gravitational potential fluctuations that are independent of  $k$ ) correspond to  $n_s = 1$ . In inflationary models [35] (see upcoming review on inflation), perturbations are generated by quantum fluctuations, which are set by the energy scale of inflation, together with the slope and higher derivatives of the inflationary potential. One generally expects that the Taylor series expansion of  $\ln \mathcal{R}_k(\ln k)$  has terms of steadily decreasing size. For the simplest models, there are thus two parameters describing the initial conditions for density perturbations, namely the amplitude and slope of the power spectrum. These can be explicitly defined, for example, through:

$$\Delta_{\mathcal{R}}^2 \equiv (k^3/2\pi^2) \langle |\mathcal{R}|_k^2 \rangle \simeq A_s (k/k_0)^{n_s-1},$$

with  $A_s \equiv \Delta_{\mathcal{R}}^2(k_0)$  and  $k_0 = 0.05 \text{ Mpc}^{-1}$ , say. There are many other equally valid definitions of the amplitude parameter (see also Secs. 22, 25, and upcoming review on inflation of this *Review*), and we caution that the relationships between some of them can be cosmology-dependent. In ‘slow roll’ inflationary models, this normalization is proportional to the combination  $V^3/(V')^2$ , for the inflationary potential  $V(\phi)$ . The slope  $n_s$  also involves  $V''$ , and so the combination of  $A_s$  and  $n_s$  can constrain potentials.

Inflation generates tensor (gravitational wave) modes, as well as scalar (density perturbation) modes. This fact introduces another parameter, measuring the amplitude of a possible tensor component, or equivalently the ratio of the tensor to scalar contributions. The tensor amplitude is  $A_t \propto V$ , and thus one expects a larger gravitational wave contribution in models where inflation happens at higher energies. The tensor power spectrum also has a slope, often denoted  $n_t$ , but since this seems unlikely to be measured in the near future, it is sufficient for now to focus only on the amplitude of the gravitational wave component. It is most common to define the tensor contribution through  $r$ , the ratio of tensor to scalar perturbation spectra at some small value of  $k$  (although sometimes it is defined in terms of the ratio of contributions at  $\ell = 2$ ). Different inflationary potentials will lead to different predictions, *e.g.*, for 50 e-folds  $\lambda\phi^4$  inflation gives  $r = 0.32$  and  $m^2\phi^2$  inflation gives  $r = 0.16$  (both now disfavored by the data), while other models can have arbitrarily small values of  $r$ . In any case, whatever the specific definition, and whether they come from inflation or something else, the ‘initial conditions’ give rise to a minimum of three parameters,  $A_s$ ,  $n_s$ , and  $r$ .

#### 28.4.2. Background Cosmology Parameters :

The FRW cosmology requires an expansion parameter (the Hubble Constant,  $H_0$ , often represented through  $H_0 = 100 h \text{ km s}^{-1} \text{ Mpc}^{-1}$ ) and several parameters to describe the matter and energy content of the Universe. These are usually given in terms of the critical density, *i.e.*, for species ‘x,’  $\Omega_x \equiv \rho_x/\rho_{\text{crit}}$ , where  $\rho_{\text{crit}} \equiv 3H_0^2/8\pi G$ . Since physical densities  $\rho_x \propto \Omega_x h^2 \equiv \omega_x$  are what govern the physics of the CMB anisotropies, it is these  $\omega$ s that are best constrained by CMB data. In particular CMB, observations constrain  $\Omega_b h^2$  for baryons and  $\Omega_c h^2$  for cold dark matter (with  $\rho_m = \rho_c + \rho_b$  for the sum).

The contribution of a cosmological constant  $\Lambda$  (or other form of dark energy, see Dark Energy—Sec. 27) is usually included via a parameter that quantifies the curvature,  $\Omega_K \equiv 1 - \Omega_{\text{tot}}$ , where  $\Omega_{\text{tot}} = \Omega_m + \Omega_\Lambda$ . The radiation content, while in principle a free parameter, is precisely enough determined by the measurement of  $T_\gamma$ , and makes a  $< 10^{-4}$  contribution to  $\Omega_{\text{tot}}$  today.

Astrophysical processes at relatively low redshift can also affect the  $C_\ell$ s, with a particularly significant effect coming through reionization. The Universe became reionized at some redshift  $z_i$ , long after recombination, affecting the CMB through the integrated Thomson scattering optical depth:

$$\tau = \int_0^{z_i} \sigma_T n_e(z) \frac{dt}{dz} dz,$$

where  $\sigma_T$  is the Thomson cross-section,  $n_e(z)$  is the number density of free electrons (which depends on astrophysics), and  $dt/dz$  is fixed by the background cosmology. In principle,  $\tau$  can be determined from the small-scale matter power spectrum, together with the physics of structure formation and radiative feedback processes; however, this is a sufficiently intricate calculation that in practice  $\tau$  needs to be considered as a free parameter.

Thus, we have eight basic cosmological parameters:  $A_s$ ,  $n_s$ ,  $r$ ,  $h$ ,  $\Omega_b h^2$ ,  $\Omega_c h^2$ ,  $\Omega_{\text{tot}}$ , and  $\tau$ . One can add additional parameters to this list, particularly when using the CMB in combination with other data sets. The next most relevant ones might be:  $\Omega_\nu h^2$ , the massive neutrino contribution;  $w$  ( $\equiv p/\rho$ ), the equation of state parameter for the dark energy; and  $dn_s/d \ln k$ , measuring deviations from a constant spectral index. To these 11 one could of course add further parameters describing additional physics, such as details of the reionization process, features in the initial power spectrum, a sub-dominant contribution of isocurvature modes, etc.

As well as these underlying parameters, there are other (dependent) quantities that can be obtained from them. Such derived parameters include the actual  $\Omega$ s of the various components (e.g.,  $\Omega_m$ ), the variance of density perturbations at particular scales (e.g.,  $\sigma_8$ ), the angular scale of the sound horizon ( $\theta_*$ ), the age of the Universe today ( $t_0$ ), the age of the Universe at recombination, reionization, etc. (see The Cosmological Parameters—Sec. 25).

## 28.5. Physics of Anisotropies

The cosmological parameters affect the anisotropies through the well understood physics of the evolution of linear perturbations within a background FRW cosmology. There are very effective, fast, and publicly available software codes for computing the CMB anisotropy, polarization, and matter power spectra, e.g., CMBFAST [36] and CAMB [37]. These have been tested over a wide range of cosmological parameters and are considered to be accurate to much better than the 1% level [38], so that numerical errors are less than 10% of the parameter uncertainties for *Planck* [9].

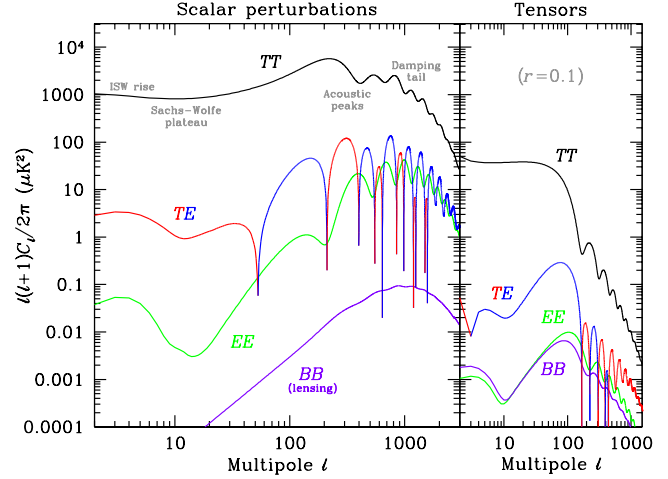
For pedagogical purposes, it is easiest to focus on the temperature anisotropies, before moving to the polarization power spectra. A description of the physics underlying the  $C_\ell^{TT}$ s can be separated into four main regions (the first two combined below), as shown in the top left part of Fig. 28.1.

### 28.5.1. The ISW Rise, $\ell \lesssim 10$ , and Sachs-Wolfe Plateau, $10 \lesssim \ell \lesssim 100$ :

The horizon scale (or more precisely, the angle subtended by the Hubble radius) at last scattering corresponds to  $\ell \simeq 100$ . Anisotropies at larger scales have not evolved significantly, and hence directly reflect the ‘initial conditions.’ Temperature variations are  $\delta T/T = -(1/5)\mathcal{R}(\mathbf{x}_{\text{LSS}}) \simeq (1/3)\delta\phi/c^2$ , where  $\delta\phi$  is the perturbation to the gravitational potential, evaluated on the last scattering surface (LSS). This is a result of the combination of gravitational redshift and intrinsic temperature fluctuations, and is usually referred to as the Sachs-Wolfe effect [39].

Assuming that a nearly scale-invariant spectrum of curvature and corresponding density perturbations was laid down at early times (i.e.,  $n_s \simeq 1$ , meaning equal power per decade in  $k$ ), then  $\ell(\ell+1)C_\ell \simeq \text{constant}$  at low  $\ell$ s. This effect is hard to see unless the multipole axis is plotted logarithmically (as in Fig. 28.1, and part of Fig. 28.2).

Time variation of the potentials (i.e., time-dependent metric perturbations) leads to an upturn in the  $C_\ell$ s in the lowest several multipoles; any deviation from a total equation of state  $w = 0$  has such an effect. So the dominance of the dark energy at low redshift (see Dark Energy—Sec. 27) makes the lowest  $\ell$ s rise above the plateau. This is sometimes called the integrated Sachs-Wolfe



**Figure 28.1:** Theoretical CMB anisotropy power spectra, using the best-fitting  $\Lambda$ CDM model from *Planck*, calculated using CAMB. The panel on the left shows the theoretical expectation for scalar perturbations, while the panel on the right is for tensor perturbations with an amplitude set to  $r = 0.1$ . Note that the  $x$ -axis is logarithmic here. For the well-measured scalar  $TT$  spectrum, the regions, each covering roughly a decade in  $\ell$ , are labeled as in the text: the ISW rise; Sachs-Wolfe plateau; acoustic peaks; and damping tail. The  $TE$  cross-correlation power spectra change sign, and that has been indicated by plotting the absolute value, but switching color for the negative parts.

effect (or ISW rise), since it comes from the line integral of  $\dot{\phi}$ ; it has been confirmed through correlations between the large-angle anisotropies and large-scale structure [40]. Specific models can also give additional contributions at low  $\ell$  (e.g., perturbations in the dark energy component itself [41]), but typically these are buried in the cosmic variance.

In principle, the mechanism that produces primordial perturbations could generate scalar, vector, and tensor modes. However, the vector (vorticity) modes decay with the expansion of the Universe. The tensors (transverse trace-free perturbations to the metric) generate temperature anisotropies through the integrated effect of the locally anisotropic expansion of space. Since the tensor modes also redshift away after they enter the horizon, they contribute only to angular scales above about  $1^\circ$  (see Fig. 28.1). Hence some fraction of the low- $\ell$  signal could be due to a gravitational wave contribution, although small amounts of tensors are essentially impossible to discriminate from other effects that might raise the level of the plateau. Nevertheless, the tensors *can* be distinguished using polarization information (see Sec. 28.7).

### 28.5.2. The Acoustic Peaks, $100 \lesssim \ell \lesssim 1000$ :

On sub-degree scales, the rich structure in the anisotropy spectrum is the consequence of gravity-driven acoustic oscillations occurring before the atoms in the Universe became neutral. Perturbations inside the horizon at last scattering have been able to evolve causally and produce anisotropy at the last scattering epoch, which reflects this evolution. The frozen-in phases of these sound waves imprint a dependence on the cosmological parameters, which gives CMB anisotropies their great constraining power.

The underlying physics can be understood as follows. Before the Universe became neutral, the proton-electron plasma was tightly coupled to the photons, and these components behaved as a single ‘photon-baryon fluid.’ Perturbations in the gravitational potential, dominated by the dark matter component, were steadily evolving. They drove oscillations in the photon-baryon fluid, with photon pressure providing most of the restoring force and baryons giving some additional inertia. The perturbations were quite small in amplitude,  $O(10^{-5})$ , and so evolved linearly. That means each Fourier mode developed independently, and hence can be described by a driven harmonic oscillator, with frequency determined by the sound speed in

the fluid. Thus the fluid density underwent oscillations, giving time variations in temperature. These combine with a velocity effect, which is  $\pi/2$  out of phase and has its amplitude reduced by the sound speed.

After the Universe recombined, the radiation decoupled from the baryons and could travel freely towards us. At that point, the (temporal) phases of the oscillations were frozen-in, and became projected on the sky as a harmonic series of peaks. The main peak is the mode that went through 1/4 of a period, reaching maximal compression. The even peaks are maximal *under*-densities, which are generally of smaller amplitude because the rebound has to fight against the baryon inertia. The troughs, which do not extend to zero power, are partially filled by the Doppler effect because they are at the velocity maxima.

The physical length scale associated with the peaks is the sound horizon at last scattering, which can be straightforwardly calculated. This length is projected onto the sky, leading to an angular scale that depends on the geometry of space, as well as the distance to last scattering. Hence the angular position of the peaks is a sensitive probe of a particular combination of cosmological parameters. In fact, the angular scale,  $\theta_s$ , is the most precisely measured observable, and hence is often treated as an element of the cosmological parameter set.

One additional effect arises from reionization at redshift  $z_i$ . A fraction of photons ( $\tau$ ) will be isotropically scattered at  $z < z_i$ , partially erasing the anisotropies at angular scales smaller than those subtended by the Hubble radius at  $z_i$ . This corresponds typically to  $\ell$ s above a few 10s, depending on the specific reionization model. The acoustic peaks are therefore reduced by a factor  $e^{-2\tau}$  relative to the plateau.

These peaks were a clear theoretical prediction going back to about 1970 [42]. One can think of them as a snapshot of stochastic standing waves. Since the physics governing them is simple and their structure rich, then one can see how they encode extractable information about the cosmological parameters. Their empirical existence started to become clear around 1994 [43], and the emergence, over the following decade, of a coherent series of acoustic peaks and troughs is a triumph of modern cosmology. This picture has received further confirmation with the detection in the power spectrum of galaxies (at redshifts close to zero) of the imprint of these same acoustic oscillations in the baryon component [44], as well as through detection of the expected oscillations in CMB polarization power spectra (see Sec. 28.7).

### 28.5.3. The Damping Tail, $\ell \gtrsim 1000$ :

The recombination process is not instantaneous, which imparts a thickness to the last scattering surface. This leads to a damping of the anisotropies at the highest  $\ell$ s, corresponding to scales smaller than that subtended by this thickness. One can also think of the photon-baryon fluid as having imperfect coupling, so that there is diffusion between the two components, and hence the amplitudes of the oscillations decrease with time. These effects lead to a damping of the  $C_\ell$ s, sometimes called Silk damping [45], which cuts off the anisotropies at multipoles above about 2000. So, although in principle it is possible to measure to ever smaller scales, this becomes increasingly difficult in practice.

### 28.5.4. Gravitational Lensing Effects :

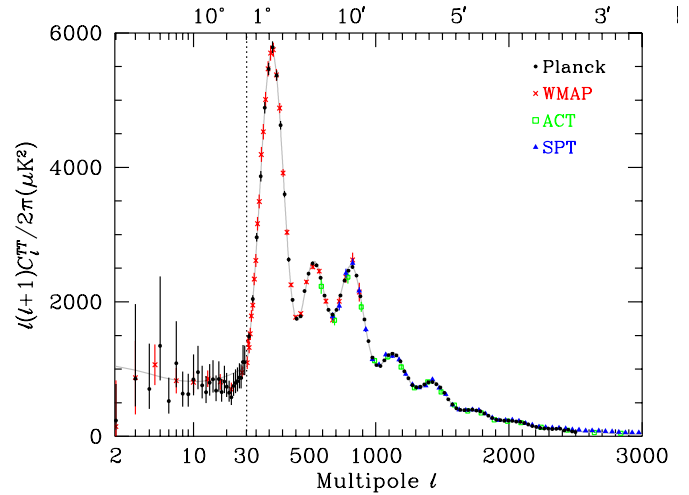
An extra effect at high  $\ell$ s comes from gravitational lensing, caused mainly by non-linear structures at low redshift. The  $C_\ell$ s are convolved with a smoothing function in a calculable way, partially flattening the peaks and troughs, generating a power-law tail at the highest multipoles, and complicating the polarization signal [46]. The effects of lensing on the CMB have now been definitively detected through the 4-point function, which correlates temperature gradients and small-scale anisotropies (enabling a map of the lensing potential to be constructed [47,48]), as well as through the smoothing effect on the shape of the  $C_\ell$ s. Lensing is important because it gives an independent estimate of  $A_s$ , breaking the parameter combination  $A_s e^{-2\tau}$  that is largely degenerate in the anisotropy power spectra.

Lensing is an example of a ‘secondary effect,’ *i.e.*, the processing of anisotropies due to relatively nearby structures (see Sec. 28.8.2). Galaxies and clusters of galaxies give several such effects; all are expected to be of low amplitude, but are increasingly important at the highest  $\ell$ s. Such effects carry additional cosmological information (about evolving gravitational potentials in the low-redshift Universe)

and are increasing in importance as experiments push to higher sensitivity and angular resolution. Measurements of the lensing power spectrum at high  $\ell$  are a particularly sensitive probe of the sum of the neutrino masses [49].

## 28.6. Current Temperature Anisotropy Data

There has been a steady improvement in the quality of CMB data that has led to the development of the present-day cosmological model. The most robust constraints currently available come from *Planck* satellite [50,51] data combined with smaller scale results from the ACT [52] and SPT [53] experiments (together with constraints from non-CMB cosmological data-sets). We plot power spectrum estimates from these experiments in Fig. 28.2, along with *WMAP* data [8] to show the consistency (see previous versions of this review for data from earlier experiments). Comparisons among data-sets show very good agreement, both in maps and in derived power spectra (up to systematic uncertainties in the overall calibration for some experiments). This makes it clear that systematic effects are largely under control.



**Figure 28.2:** CMB temperature anisotropy band-power estimates from the *Planck*, *WMAP*, ACT, and SPT experiments. Note that the widths of the  $\ell$ -bands vary between experiments and have not been plotted. This figure represents only a selection of the most recent available experimental results, and some points with large error bars have been omitted. At the higher multipoles these band-powers involve subtraction of particular foreground models, while proper analysis requires simultaneous fitting of CMB and foregrounds over multiple frequencies. The  $x$ -axis here is logarithmic for the lowest multipoles, to show the Sachs-Wolfe plateau, and linear for the other multipoles. The acoustic peaks and damping region are very clearly observed, with no need for a theoretical curve to guide the eye; however, the curve plotted is the best-fit *Planck* model.

The band-powers shown in Fig. 28.2 are in very good agreement with a ‘ $\Lambda$ CDM’ model. As described earlier, several (at least eight) of the peaks and troughs are quite apparent. For details of how these estimates were arrived at, the strength of correlations between band-powers and other information required to properly interpret them, the original papers should be consulted.

## 28.7. CMB Polarization

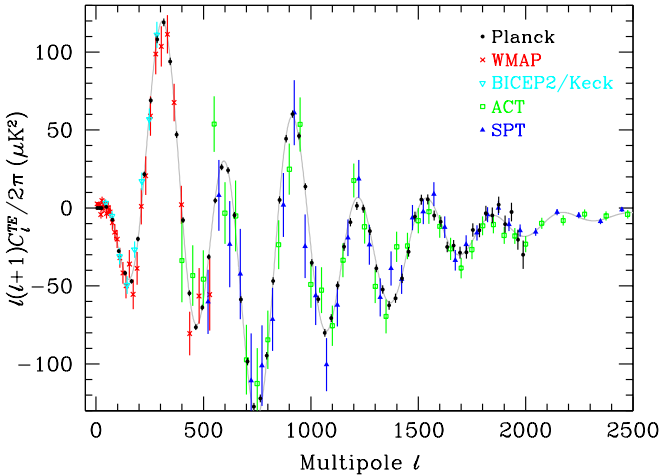
Since Thomson scattering of an anisotropic radiation field also generates linear polarization, the CMB is predicted to be polarized at the level of roughly 5% of the temperature anisotropies [54]. Polarization is a spin-2 field on the sky, and the algebra of the modes in  $\ell$ -space is strongly analogous to spin-orbit coupling in quantum mechanics [55]. The linear polarization pattern can be decomposed in a number of ways, with two quantities required for each pixel in



a map, often given as the  $Q$  and  $U$  Stokes parameters. However, the most intuitive and physical decomposition is a geometrical one, splitting the polarization pattern into a part that comes from a divergence (often referred to as the ‘ $E$ -mode’) and a part with a curl (called the ‘ $B$ -mode’) [56]. More explicitly, the modes are defined in terms of second derivatives of the polarization amplitude, with the Hessian for the  $E$ -modes having principle axes in the same sense as the polarization, while the  $B$ -mode pattern can be thought of as a  $45^\circ$  rotation of the  $E$ -mode pattern. Globally one sees that the  $E$ -modes have  $(-1)^\ell$  parity (like the spherical harmonics), while the  $B$ -modes have  $(-1)^{\ell+1}$  parity.

The existence of this linear polarization allows for six different cross power spectra to be determined from data that measure the full temperature and polarization anisotropy information. Parity considerations make two of these zero, and we are left with four potential observables,  $C_\ell^{TT}$ ,  $C_\ell^{TE}$ ,  $C_\ell^{EE}$ , and  $C_\ell^{BB}$  (see Fig. 28.1). Because scalar perturbations have no handedness, the  $B$ -mode power spectrum can only be sourced by vectors or tensors. Moreover, since inflationary scalar perturbations give only  $E$ -modes, while tensors generate roughly equal amounts of  $E$ - and  $B$ -modes, then the determination of a non-zero  $B$ -mode signal is a way to measure the gravitational wave contribution (and thus potentially derive the energy scale of inflation). However, since the signal is expected to be rather weak, one must first eliminate the foreground contributions and other systematic effects down to very low levels.

The polarization  $C_\ell$ s also exhibit a series of acoustic peaks generated by the oscillating photon-baryon fluid. The main ‘ $EE$ ’ power spectrum has peaks that are out of phase with those in the ‘ $TT$ ’ spectrum, because the polarization anisotropies are sourced by the fluid velocity. The ‘ $TE$ ’ part of the polarization and temperature patterns comes from correlations between density and velocity perturbations on the last scattering surface, which can be both positive and negative, and is of larger amplitude than the  $EE$  signal. There is no polarization Sachs-Wolfe effect, and hence no large-angle plateau. However, scattering during a recent period of reionization can create a polarization ‘bump’ at large angular scales.



**Figure 28.3:** Cross-power spectrum band-powers of the temperature anisotropies and  $E$ -mode polarization signal from *Planck* (the low multipole data have been binned here), *WMAP*, *BICEP2/Keck*, *ACT*, and *SPT*. The curve is the prediction from the best fit to the *Planck* temperature band-powers (as well as the  $\ell < 30$  polarization and CMB lensing results) and is not a fit to these data; however, these  $TE$  measurements follow the curve very closely, showing the expected oscillatory structure. Note that each band-power is an average over a range of multipoles, and hence to compare in detail with a model one has to average the theoretical curve through the band.

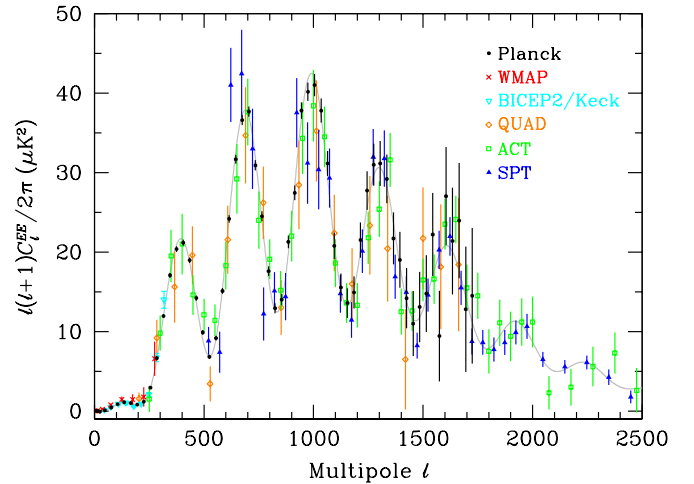
Because the polarization anisotropies have only a fraction of the amplitude of the temperature anisotropies, they took longer to detect. The first measurement of a polarization signal came in 2002 from

the DASI experiment [57], which provided a convincing detection, confirming the general paradigm, but of low enough significance that it lent little constraint to models. Despite dramatic progress since then, it is still the case that polarization data mainly support the basic paradigm, without dramatically reducing error bars on parameters. However, there are exceptions to this, specifically in the reionization optical depth, and the potential to constrain primordial gravitational waves.

### 28.7.1. $T$ - $E$ Power Spectrum :

Since the  $T$  and  $E$  skies are correlated, one has to measure the  $TE$  power spectrum, as well as  $TT$  and  $EE$ , in order to extract all the cosmological information. This  $TE$  signal has now been mapped out extremely accurately by *Planck* [51], and these band-powers are shown in Fig. 28.3, along with those from *WMAP* [58] and *BICEP2/Keck* [59], with *ACT* [60] and *SPT* [61] extending to smaller angular scales. The anti-correlation at  $\ell \simeq 150$  and the peak at  $\ell \simeq 300$  were the first features to become distinct, but now a whole series of oscillations is clearly seen in this power spectrum. The measured shape of the cross-correlation power spectrum provides supporting evidence for the general cosmological picture, as well as directly constraining the thickness of the last scattering surface. Since the polarization anisotropies are generated in this scattering surface, the existence of correlations at angles above about a degree demonstrates that there were super-Hubble fluctuations at the recombination epoch. The sign of this correlation also confirms the adiabatic paradigm.

The overall picture of the source of CMB polarization and its oscillations has also been confirmed through tests that average the maps around both temperature hot spots and cold spots [62,11]. One sees precisely the expected patterns of radial and tangential polarization configurations, as well as the phase shift between polarization and temperature. This leaves no doubt that the oscillation picture is the correct one and that the polarization is coming from Thomson scattering at  $z \simeq 1100$ .



**Figure 28.4:** Power spectrum of  $E$ -mode polarization from *Planck*, together with *WMAP*, *BICEP2/Keck*, *QUAD*, *ACT*, and *SPT*. Note that some band-powers with larger uncertainties have been omitted and that the unbinned *Planck* low- $\ell$  data have been binned here. Also plotted is the best-fit theoretical model from *Planck*  $TT$  data (plus polarization at  $\ell < 30$  and CMB lensing).

### 28.7.2. $E$ - $E$ Power Spectrum :

Experimental band-powers for  $C_\ell^{EE}$  from *Planck*, *WMAP*, *BICEP2/Keck* [59], *QUAD* [63], *ACT* [60], and *SPT* [61] are shown in Fig. 28.4. Without the benefit of correlating with the temperature anisotropies (i.e., measuring  $C_\ell^{TE}$ ), the polarization anisotropies are very weak and challenging to measure. Nevertheless, the oscillatory pattern is becoming well established and the data closely match the  $TT$ -derived theoretical prediction. In Fig. 28.4 one can clearly see the ‘shoulder’ expected at  $\ell \simeq 140$ , the first main peak

at  $\ell \simeq 400$  (corresponding to the first trough in  $C_\ell^{TT}$ ), and the series of oscillations that is out of phase with the temperature anisotropy power spectrum.

Perhaps the most distinctive result from the polarization measurements is at the largest angular scales ( $\ell < 10$ ) in  $C_\ell^{TE}$  and  $C_\ell^{EE}$ , where there is evidence for an excess signal (not visible in Fig. 28.4) compared to that expected from the temperature power spectrum alone. This is precisely the signal anticipated from an early period of reionization, arising from Doppler shifts during the partial scattering at  $z < z_i$ . The amplitude of the signal indicates that the first stars, presumably the source of the ionizing radiation, formed around  $z \simeq 10$  (although the uncertainty is still quite large). Since this corresponds to scattering optical depth  $\tau \simeq 0.1$ , then roughly 10% of CMB photons were re-scattered at the reionization epoch, with the other 90% last scattering at  $z \simeq 1100$ . However, estimates of the amplitude of this reionization excess have come down since the first measurements by *WMAP* (indicating that this is an extremely difficult measurement to make) and the latest determination from *Planck* gives  $z_i = 8.8^{+1.7}_{-1.4}$  [13].

### 28.7.3. *B-B Power Spectrum* :

The expected amplitude of  $C_\ell^{BB}$  is very small, and so measurements of this polarization curl-mode are very challenging. The first indication of the existence of the *BB* signal has come from the detection of the expected conversion of *E*-modes to *B*-modes by gravitational lensing, through a correlation technique using the lensing potential and polarization measurements from SPT [64]. However, the real promise of *B*-modes lies in the detection of primordial gravitational waves at larger scales. This tensor signature could be seen either in the ‘recombination bump’ at around  $\ell = 100$  (caused by an ISW effect as gravitational waves redshift away at the last-scattering epoch) or the ‘reionization bump’ (from additional scattering at low redshifts).

Results from the BICEP-2 experiment [65] in 2014 suggested a detection of the primordial *B*-mode signature around the recombination peak. BICEP-2 mapped a small part of the CMB sky with the the lowest sensitivity level yet reached (below 100 nK), but at a single frequency. Higher frequency data from *Planck* indicated that much of the BICEP2 signal was due to dust within our Galaxy, and a combined analysis by the BICEP-2, Keck Array, and *Planck* teams [66] indicated that the data are consistent with no primordial *B*-modes, with an upper limit of  $r < 0.12$ .

Several experiments are continuing to push down the sensitivity of *B*-mode measurements, motivated by the enormous importance of a future detection of this telltale signature of inflation (or other physics at the highest energies). A compilation of experimental results for  $C_\ell^{BB}$  is shown in Fig. 28.5, coming from a combination of direct estimates of the *B*-modes (BICEP2/Keck Array [59], POLARBEAR [67], and SPTpol [68]) and indirect determinations of the lensing *B*-modes based on estimating the effect of measured lensing on measured *E*-modes (*Planck* [48], SPT [64], and ACT [69]). Additional band-power estimates are expected from these and other experiments in the near future.

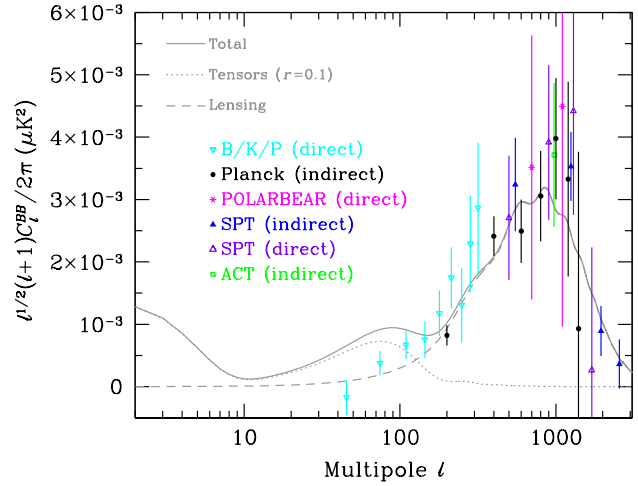
## 28.8. Complications

There are a number of issues that complicate the interpretation of CMB anisotropy data (and are considered to be *signal* by many astrophysicists), some of which we sketch out below.

### 28.8.1. *Foregrounds* :

The microwave sky contains significant emission from our Galaxy and from extra-galactic sources [70]. Fortunately, the frequency dependence of these various sources is in general substantially different from that of the CMB anisotropy signals. The combination of Galactic synchrotron, bremsstrahlung, and dust emission reaches a minimum at a frequency of roughly 100 GHz (or wavelength of about 3 mm). As one moves to greater angular resolution, the minimum moves to slightly higher frequencies, but becomes more sensitive to unresolved (point-like) sources.

At frequencies around 100 GHz, and for portions of the sky away from the Galactic Plane, the foregrounds are typically 1 to 10% of the CMB anisotropies. By making observations at multiple frequencies, it is relatively straightforward to separate the various components and determine the CMB signal to the few per cent level.



**Figure 28.5:** Power spectrum of *B*-mode polarization, including results from the BICEP2/Keck Array/*Planck* combined analysis (B/K/P), *Planck*, POLARBEAR, SPT, and ACT. Note that some of the measurements are direct estimates of *B*-modes on the sky, while others are only sensitive to the lensing signal and come from combining *E*-mode and lensing potential measurements. Several experiments have previously reported upper limits, which are all off the top of this plot. A logarithmic  $x$ -axis is adopted here and the  $y$ -axis has been divided by a factor of  $\sqrt{\ell}$  in order to show all three theoretically expected contributions: the low- $\ell$  reionization bump; the  $\ell \sim 100$  recombination peak; and the high- $\ell$  lensing signature. The dotted line is for a tensor (primordial gravitational wave) fraction  $r = 0.1$ , simply as an example, with all other cosmological parameters set at the best *Planck*-derived values, for which model the expected lensing *B*-modes have also been shown with a dashed line.

For greater sensitivity, it is necessary to use the spatial information and statistical properties of the foregrounds to separate them from the CMB. Furthermore, at higher  $\ell$ s it is essential to carefully model extragalactic foregrounds, particularly the clustering of infrared-emitting galaxies, which dominate the measured power spectrum as we move into the damping tail.

The foregrounds for CMB polarization follow a similar pattern to those for temperature, but are less well studied, and are intrinsically brighter relative to CMB anisotropies. *WMAP* showed that the polarized foregrounds dominate at large angular scales, and that they must be well characterized in order to be discriminated [71]. *Planck* has shown that it is possible to characterize the foreground polarization signals, with synchrotron dominating at low frequencies and dust at high frequencies [72]. Whether the foregrounds become more complicated as we push down in sensitivity at high multipoles is not known. However, although they make analysis more difficult, for the time being, foreground contamination is not a fundamental limit for CMB experiments.

### 28.8.2. *Secondary Anisotropies* :

With increasingly precise measurements of the primary anisotropies, there is growing theoretical and experimental interest in ‘secondary anisotropies,’ pushing experiments to higher angular resolution and sensitivity. These secondary effects arise from the processing of the CMB due to ionization history and the evolution of structure, including gravitational lensing (which was already discussed) and patchy reionization effects [73]. Additional information can thus be extracted about the Universe at  $z \ll 1000$ . This tends to be most effectively done through correlating CMB maps with other cosmological probes of structure. Secondary signals are also typically non-Gaussian, unlike the primary CMB anisotropies.

A secondary signal of great current interest is the Sunyaev-Zeldovich (SZ) effect [74], which is Compton scattering ( $\gamma e \rightarrow \gamma' e'$ ) of the CMB photons by hot electron gas. This creates spectral distortions by transferring energy from the electrons to the photons. It is particularly important for clusters of galaxies, through which one



observes a partially Comptonized spectrum, resulting in a decrement at radio wavelengths and an increment in the submillimeter.

The imprint on the CMB sky is of the form  $\Delta T/T = y f(x)$ , with the  $y$ -parameter being the integral of Thomson optical depth times  $kT_e/m_e c^2$  through the cluster, and  $f(x)$  describing the frequency dependence. This is simply  $x \coth(x/2) - 4$  for a non-relativistic gas (the electron temperature in a cluster is typically a few keV), where the dimensionless frequency  $x \equiv h\nu/kT_\gamma$ . As well as this ‘thermal’ SZ effect, there is also a smaller ‘kinetic’ effect due to the bulk motion of the cluster gas, giving  $\Delta T/T \sim \tau(v/c)$ , with either sign, but having the same spectrum as the primary CMB anisotropies.

A significant advantage in finding galaxy clusters this way is that the SZ effect is largely independent of redshift, so in principle clusters can be found to arbitrarily large distances. The SZ effect can be used to find and study individual clusters, and to obtain estimates of the Hubble constant. There is also the potential to constrain cosmological parameters, such as the clustering amplitude  $\sigma_8$  and the equation of state of the dark energy, through counts of detected clusters as a function of redshift. The promise of the method has been realized through detections of clusters purely through the SZ effect by SPT [75], ACT [76], and *Planck* [77]. Results from *Planck* clusters [78] suggest a somewhat lower value of  $\sigma_8$  than inferred from CMB anisotropies, but there are still systematic uncertainties that might encompass the difference. Further analysis of scaling relations among cluster properties should enable more robust cosmological constraints to be placed in future, so that we can understand whether this ‘tension’ might be a sign of new physics.

### 28.8.3. Higher-order Statistics :

Although most of the CMB anisotropy information is contained in the power spectra, there will also be weak signals present in higher-order statistics. These can measure any primordial non-Gaussianity in the perturbations, as well as non-linear growth of the fluctuations on small scales and other secondary effects (plus residual foreground contamination of course). Although there are an infinite variety of ways in which the CMB could be non-Gaussian [29], there is a generic form to consider for the initial conditions, where a quadratic contribution to the curvature perturbations is parameterized through a dimensionless number  $f_{\text{NL}}$ . This weakly non-linear component can be constrained in several ways, the most popular being through measurements of the bispectrum.

The constraints depend on the shape of the triangles in harmonic space, and it has become common to distinguish the ‘local’ or ‘squeezed’ configuration (in which one side is much smaller than the other two) from the ‘equilateral’ configuration. Other configurations are also relevant for specific theories, such as ‘orthogonal’ non-Gaussianity, which has positive correlations for  $k_1 \simeq 2k_2 \simeq 2k_3$ , and negative correlations for the equilateral configuration. The results from the *Planck* team [79] (including polarization here) are  $f_{\text{NL}}^{\text{local}} = 1 \pm 5$ ,  $f_{\text{NL}}^{\text{equil}} = 0 \pm 40$ , and  $f_{\text{NL}}^{\text{ortho}} = -26 \pm 21$ .

These results are consistent with zero, but are at a level that is now interesting for model predictions. The amplitude of  $f_{\text{NL}}$  expected is small, so that a detection of  $f_{\text{NL}} \gg 1$  would rule out all single-field, slow-roll inflationary models. It is still possible to improve upon these *Planck* results, and it certainly seems feasible that a measurement of primordial non-Gaussianity may yet be within reach. Non-primordial detections of non-Gaussianity from expected signatures have already been made. For example, the bispectrum and trispectrum contain evidence of gravitational lensing, the ISW effect, and Doppler boosting. For now the primordial signal is elusive, but should it be detected, then detailed measurements of non-Gaussianity will become a unique probe of inflationary-era physics. Because of that, much effort continues to be devoted to honing predictions and measurement techniques, with the expectation that we will need to go beyond the CMB to dramatically improve the constraints.

### 28.8.4. Anomalies :

Several features seen in the *Planck* data [32] confirm those found earlier with *WMAP* [31], showing mild deviations from a simple description of the data; these are often referred to as ‘anomalies.’ One such feature is the apparent lack of power in the multipole range  $\ell \simeq 20\text{--}30$  [9,51]. The other examples involve the breaking of statistical isotropy, caused by alignment of the lowest multipoles,

or a somewhat excessive cold spot, or a power asymmetry between hemispheres. No such feature is significant at more than the roughly  $3\sigma$  level, and the importance of ‘a posteriori’ statistics here has been emphasized by many authors. Since these effects are at large angular scales, where cosmic variance dominates, the results will not increase in significance with more data, although there is the potential for polarization data to provide independent tests.

## 28.9. Constraints on Cosmological Parameters

The most striking outcome of the newer experimental results is that the standard cosmological paradigm is in very good shape. A large amount of high precision data on the power spectrum is adequately fit with fewer than 10 free parameters (and only six need non-trivial values). The framework is that of FRW models, which have nearly flat geometry, containing dark matter and dark energy, and with adiabatic perturbations having close to scale-invariant initial conditions.

Within this basic picture, the values of the cosmological parameters can be constrained. Of course, much more stringent bounds can be placed on models that cover a restricted parameter space, *e.g.*, assuming that  $\Omega_{\text{tot}} = 1$  or  $r = 0$ . More generally, the constraints depend upon the adopted prior probability distributions, even if they are implicit, for example by restricting the parameter freedom or their ranges (particularly where likelihoods peak near the boundaries), or by using different choices of other data in combination with the CMB. As the data become even more precise, these considerations will be less important, but for now we caution that restrictions on model space and choice of non-CMB data-sets and priors need to be kept in mind when adopting specific parameter values and uncertainties.

There are some combinations of parameters that fit the CMB anisotropies almost equivalently. For example, there is a nearly exact geometric degeneracy, where any combination of  $\Omega_m$  and  $\Omega_\Lambda$  that gives the same angular diameter distance to last scattering will give nearly identical  $C_\ell$ s. There are also other less exact degeneracies among the parameters. Such degeneracies can be broken when using the CMB results in combination with other cosmological data-sets. Particularly useful are complementary constraints from baryon acoustic oscillations, galaxy clustering, the abundance of galaxy clusters, weak gravitational lensing measurements, and Type Ia supernova distances. For an overview of some of these other cosmological constraints, see The Cosmological Parameters—Sec. 25 of this *Review*.

Within the context of a six parameter family of models (which fixes  $\Omega_{\text{tot}} = 1$ ,  $dn_s/d\ln k = 0$ ,  $r = 0$ , and  $w = -1$ ) the *Planck* results for  $TT$ , together with low- $\ell$  polarization and CMB lensing, and the use of high- $\ell$  data from ACT and SPT to constrain foregrounds, yields [13]:  $\ln(10^{10} A_s) = 3.062 \pm 0.029$ ;  $n_s = 0.968 \pm 0.006$ ;  $\Omega_b h^2 = 0.02226 \pm 0.00023$ ;  $\Omega_c h^2 = 0.1186 \pm 0.0020$ ;  $100\theta_* = 1.0410 \pm 0.0005$ ; and  $\tau = 0.066 \pm 0.016$ . Other parameters can be derived from this basic set, including  $h = 0.678 \pm 0.009$ ,  $\Omega_\Lambda = 0.692 \pm 0.012 (= 1 - \Omega_m)$  and  $\sigma_8 = 0.815 \pm 0.009$ . Somewhat different (although consistent) values are obtained using other data combinations, such as including BAO, supernova,  $H_0$ , or weak lensing constraints (see Sec. 25 of this *Review*). However, the results quoted above are currently the best available from CMB anisotropies alone. The uncertainties decrease by around 25% when adding *Planck* polarization data, although the recommendation for now is not to include these 2015 polarization data in fits, since there are still some unmodelled systematic effects present [51].

The standard cosmological model continues to fit the data well, with the error bars on the parameters continuing to shrink. Improved measurement of higher acoustic peaks has dramatically reduced the uncertainty in the  $\theta_*$  parameter, which is now detected at  $> 2000\sigma$ . The evidence for  $n_s < 1$  remains above the  $5\sigma$  level. The value of the reionization optical depth has decreased compared with earlier estimates; it is convincingly detected, but still not of very high significance.

Constraints can also be placed on parameters beyond the basic six, particularly when including other astrophysical data-sets. Relaxing the flatness assumption, the constraint on  $\Omega_{\text{tot}}$  is  $1.005 \pm 0.008$ . Note that for  $h$ , the CMB data alone provide only a very weak constraint if spatial flatness is not assumed. However, with the addition of other data (particularly powerful in this context being a compilation of BAO measurements [80]), the constraints on the

Hubble constant and curvature improve considerably, leading to  $\Omega_{\text{tot}} = 1.0002 \pm 0.0026$  [13].

For  $\Omega_b h^2$  the CMB-derived value is generally consistent with completely independent constraints from Big Bang nucleosynthesis (see Sec. 24 of this *Review*). Related are constraints on additional neutrino-like relativistic degrees of freedom, which lead to  $N_{\text{eff}} = 3.15 \pm 0.23$  (including BAO), *i.e.*, no evidence for extra neutrino species.

The 95% confidence upper limit on  $r$  (measured at  $k = 0.002 \text{ Mpc}^{-1}$ ) from the effect of tensors solely on  $C_\ell^{TT}$  (see Fig. 28.1) is 0.11. This limit depends on how the slope  $n$  is restricted and whether  $dn_s/d\ln k \neq 0$  is allowed. The joint constraints on  $n_s$  and  $r$  allows specific inflationary models to be tested [33,34]. The limit on  $r$  is even tighter when combined with the BICEP/Keck/*Planck* results for  $C_\ell^{BB}$ , yielding  $r < 0.08$  at 95% confidence [34]. Looking at the  $(n_s, r)$  plane, this means that  $m^2\phi^2$  (mass-term quadratic) inflation is now disfavored by the data, as well as  $\lambda\phi^4$  (self-coupled) inflation.

The addition of the dark energy equation of state  $w$  adds the partial degeneracy of being able to fit a ridge in  $(w, h)$  space, extending to low values of both parameters. This degeneracy is broken when the CMB is used in combination with other data-sets, *e.g.*, adding a compilation of BAO data gives  $w = -1.01 \pm 0.05$ . Constraints can also be placed on more general dark energy and modified gravity models [81]. However, one needs to be careful not to over-interpret some tensions between data-sets as evidence for new physics.

For the optical depth  $\tau$ , the best-fit corresponds to a reionization redshift centered on 9 in the best-fit cosmology, and assuming instantaneous reionization. This redshift is only slightly higher than that suggested from studies of absorption lines in high- $z$  quasar spectra [82] and Ly $\alpha$ -emitting galaxies [83], perhaps hinting that the process of reionization was not as complex as previously suspected. The important constraint provided by CMB polarization, in combination with astrophysical measurements, thus allows us to investigate how the first stars formed and brought about the end of the cosmic dark ages.

## 28.10. Particle Physics Constraints

CMB data place limits on parameters that are directly relevant for particle physics models. For example, there is a limit on the sum of the masses of the neutrinos,  $\sum m_\nu < 0.21 \text{ eV}$  (95%) [9] coming from *Planck* together with BAO measurements (although limits are weaker when considering both  $N_{\text{eff}}$  and  $\sum m_\nu$  as free parameters). This assumes the usual number density of fermions, which decoupled when they were relativistic. The limit is tantalizingly only a factor of a few higher than the minimum value coming from neutrino mixing experiments (see Neutrino Mixings—Sec. 14). As well as being an indirect probe of the neutrino background, *Planck* data also require that the neutrino background has perturbations, *i.e.* that it possesses a sound speed  $c_s^2 \simeq 1/3$ , as expected [13].

The current suite of data suggests that  $n_s < 1$ , with a best-fitting value about 0.03 below unity. This is already quite constraining for inflationary models, particularly along with  $r$  limits. There is no current evidence for running of the spectral index, with  $dn_s/d\ln k = -0.003 \pm 0.008$  from *Planck* alone [13], although this is less of a constraint on inflationary models. Similarly, primordial non-Gaussianity is being probed to interesting levels, although tests of simple inflationary models will only come with significant reductions in uncertainty.

The large-angle anomalies, such as the hemispheric modulation of power and the dip in power at  $\ell \simeq 20\text{--}30$ , have the potential to be hints of new physics. Such effects might be expected in a universe that has a large-scale power cut-off, or anisotropy in the initial power spectrum, or is topologically non-trivial. However, cosmic variance and *a posteriori* statistics limit the significance of these anomalies, absent the existence of a model that naturally yields some of these features (and hopefully also predicting other phenomena that can be tested).

It is possible to place limits on additional areas of physics [84], for example annihilating dark matter, [13], primordial magnetic fields [85], and time variation of the fine-structure constant [86], as well as parity violation, the neutrino chemical potential, a contribution of warm dark matter, topological defects, or physics beyond general

relativity. Further particle physics constraints will follow as the anisotropy measurements increase in precision.

The CMB anisotropy measurements precisely pin down physics at the time of last-scattering, and so any change of physics can be constrained if it affects the relevant energies or timescales. Future, higher sensitivity measurements of the CMB frequency spectrum will push the constraints back to cover energy injection at much earlier times ( $\sim 1$  year). Comparison of CMB and BBN observables extend these constraints to timescales of order seconds, and energies in the MeV range. And to the extent that inflation provides an effective description of the generation of perturbations, the inflationary observables will constrain physics at GUT-type energy scales.

More generally, careful measurement of the CMB power spectra and non-Gaussianity can in principle put constraints on physics at the highest energies, including ideas of string theory, extra dimensions, colliding branes, *etc.* At the moment any calculation of predictions appears to be far from definitive. However, there is a great deal of activity on implications of string theory for the early Universe, and hence a very real chance that there might be observational implications for specific scenarios.

## 28.11. Fundamental Lessons

More important than the precise values of parameters is what we have learned about the general features that describe our observable Universe. Beyond the basic hot Big Bang picture, the CMB has taught us that:

- The Universe recombined at  $z \simeq 1100$  and started to become ionized again at  $z \simeq 10$ .
- The geometry of the Universe is close to flat.
- Both dark matter and dark energy are required.
- Gravitational instability is sufficient to grow all of the observed large structures in the Universe.
- Topological defects were not important for structure formation.
- There are ‘synchronized’ super-Hubble modes generated in the early Universe.
- The initial perturbations were predominantly adiabatic in nature.
- The perturbation spectrum has a slightly red tilt.
- The perturbations had close to Gaussian (*i.e.*, maximally random) initial conditions.

These features form the basis of the cosmological standard model,  $\Lambda$ CDM, for which it is tempting to make an analogy with the Standard Model of particle physics (see earlier Sections of this *Review*). The cosmological model is much further from any underlying ‘fundamental theory,’ which may ultimately provide the values of the parameters from first principles. Nevertheless, any genuinely complete ‘theory of everything’ must include an explanation for the values of these cosmological parameters as well as the parameters of the Standard Model of particle physics.

## 28.12. Future Directions

Given the significant progress in measuring the CMB sky, which has been instrumental in tying down the cosmological model, what can we anticipate for the future? There will be a steady improvement in the precision and confidence with which we can determine the appropriate cosmological parameters. Ground-based experiments operating at smaller angular scales will continue to place tighter constraints on the damping tail. New polarization experiments at small scales will probe further into the damping tail, without the limitation of extragalactic foregrounds. And polarization experiments at large angular scales will push down the limits on primordial  $B$ -modes.

*Planck*, the third generation CMB satellite mission, was launched in May 2009, and has produced a large number of papers, including a set of cosmological studies based on the first two full surveys of the sky (accompanied by a public release of data products) in 2013 and a further series based on analysis of the full mission data release in 2015 (eight surveys for the Low Frequency Instrument and five surveys for the High Frequency Instrument). In 2016 results are expected from a final analysis, including a comprehensive investigation of polarization.

A set of cosmological parameters is now known to percent level accuracy, and that may seem sufficient for many people. However,

we should certainly demand more of measurements that describe *the entire observable Universe!* Hence a lot of activity in the coming years will continue to focus on determining those parameters with increasing precision. This necessarily includes testing for consistency among different predictions of the cosmological Standard Model, and searching for signals that might require additional physics.

A second area of focus will be the smaller scale anisotropies and ‘secondary effects.’ There is a great deal of information about structure formation at  $z \ll 1000$  encoded in the CMB sky. This may involve higher-order statistics and cross-correlations with other large-scale structure tracers, as well as spectral signatures, with many experiments targeting the galaxy cluster SZ effect. The current status of CMB lensing is similar (in terms of total signal-to-noise) to the quality of the first CMB anisotropy measurements by *COBE*, and thus we can expect that experimental probes of lensing will improve dramatically in the coming years. All of these investigations can provide constraints on the dark energy equation of state, for example, which is a major area of focus for several future cosmological surveys at optical wavelengths. CMB lensing also promises to yield a measurement of the sum of the neutrino masses.

A third direction is increasingly sensitive searches for specific signatures of physics at the highest energies. The most promising of these may be the primordial gravitational wave signals in  $C_\ell^{BB}$ , which could be a probe of the  $\sim 10^{16}$  GeV energy range. There are several ground- and balloon-based experiments underway that are designed to search for the polarization *B*-modes. Additionally, non-Gaussianity holds the promise of constraining models beyond single-field slow-roll inflation.

Anisotropies in the CMB have proven to be the premier probe of cosmology and the early Universe. Theoretically the CMB involves well understood physics in the linear regime, and is under very good calculational control. A substantial and improving set of observational data now exists. Systematics appear to be under control and not a limiting factor. And so for the next few years we can expect an increasing amount of cosmological information to be gleaned from CMB anisotropies, with the prospect also of some genuine surprises.

## References:

1. A.A. Penzias and R. Wilson, *Astrophys. J.* **142**, 419 (1965); R.H. Dicke *et al.*, *Astrophys. J.* **142**, 414 (1965).
2. M. White, D. Scott, and J. Silk, *Ann. Rev. Astron. Astrophys.* **32**, 329 (1994); W. Hu and S. Dodelson, *Ann. Rev. Astron. Astrophys.* **40**, 171 (2002); A. Challinor and H. Peiris, *Proc. Astro. Soc. Pacific Conf. Ser.* Vol. 1132, 86 (2009).
3. G.F. Smoot *et al.*, *Astrophys. J.* **396**, L1 (1992).
4. C.L. Bennett *et al.*, *Astrophys. J. Supp.* **148**, 1 (2003).
5. N. Jarosik *et al.*, *Astrophys. J. Supp.* **170**, 263 (2007).
6. G. Hinshaw *et al.*, *Astrophys. J. Supp.* **180**, 225 (2009).
7. N. Jarosik *et al.*, *Astrophys. J. Supp.* **192**, 14 (2011).
8. G. Hinshaw *et al.*, *Astrophys. J. Supp.* **208**, 19 (2013).
9. Planck Collab. 2013 Results XVI, *Astron. & Astrophys.* **571**, A16 (2014).
10. J.A. Tauber *et al.*, *Astron. & Astrophys.* **520**, 1 (2010); Planck Collab. Early Results I, *Astron. & Astrophys.* **536**, 1 (2011).
11. Planck Collab. 2013 Results I, *Astron. & Astrophys.* **571**, A1 (2014).
12. Planck Collab. 2015 Results I, *Astron. & Astrophys.* submitted, [arXiv:1502.01582](#).
13. Planck Collab. 2015 Results XIII, *Astron. & Astrophys.* submitted, [arXiv:1502.01589](#).
14. D.S. Swetz *et al.*, *Astrophys. J. Supp.* **194**, 41 (2011).
15. J.E. Carlstrom *et al.*, *Publ. Astron. Soc. Pacific* **123**, 568 (2011).
16. P. Noterdaeme *et al.*, *Astron. & Astrophys.* **526**, L7 (2011).
17. D.J. Fixsen *et al.*, *Astrophys. J.* **734**, 5 (2011).
18. R. Subrahmanyam and R. Cowsik, *Astrophys. J.* **776**, 42 (2013).
19. A. Kogut *et al.*, *Proc. SPIE*, vol. 9143, 1E (2014).
20. R. Khatri and R.A. Sunyaev, *JCAP* **6**, 026 (2013); J. Chluba and D. Jeong, *MNRAS* **438**, 2065 (2014).
21. D.J. Fixsen, *Astrophys. J.* **707**, 916 (2009).
22. J.C. Mather *et al.*, *Astrophys. J.* **512**, 511 (1999).
23. Y. Hoffman, H.M. Courtois, and R.B. Tully, *MNRAS* **449**, 4494 (2015).
24. D.J. Fixsen *et al.*, *Astrophys. J.* **420**, 445 (1994).
25. Planck Collab. 2013 Results XXVII, *Astron. & Astrophys.* **571**, A27 (2014).
26. A. Kogut *et al.*, *Astrophys. J.* **419**, 1 (1993).
27. S. Seager, D.D. Sasselov, and D. Scott, *Astrophys. J. Supp.* **128**, 407 (2000).
28. L. Knox, *Phys. Rev.* **D52**, 4307 (1995).
29. N. Bartolo *et al.*, *Phys. Reports* **402**, 103 (2004); E. Komatsu, *Class. Quantum Grav.* **27**, 124010 (2010); A.P.S. Yadav and B.D. Wandelt, *Adv. Astron.* **2010**, 565248 (2010).
30. Planck Collab. 2013 Results XXIV, *Astron. & Astrophys.* **571**, A24 (2014).
31. C.L. Bennett *et al.*, *Astrophys. J. Supp.* **192**, 17 (2011).
32. Planck Collab. 2013 Results XXIII, *Astron. & Astrophys.* **571**, A23 (2014); Planck Collab. 2015 Results XXIII, *Astron. & Astrophys.* submitted, [arXiv:1506.07135](#).
33. Planck Collab. 2013 Results XXII, *Astron. & Astrophys.* **571**, A22 (2014).
34. Planck Collab. 2015 Results XX, *Astron. & Astrophys.* submitted, [arXiv:1502.02114](#).
35. A.R. Liddle and D.H. Lyth, *Cosmological Inflation and Large-Scale Structure*, Cambridge University Press (2000).
36. U. Seljak and M. Zaldarriaga, *Astrophys. J.* **469**, 437 (1996).
37. A. Lewis, A. Challinor, and A. Lasenby, *Astrophys. J.* **538**, 473 (2000).
38. U. Seljak *et al.*, *Phys. Rev.* **D68**, 083507 (2003); C. Howlett *et al.*, *JCAP* **04**, 027 (2012).
39. R.K. Sachs and A.M. Wolfe, *Astrophys. J.* **147**, 73 (1967).
40. R. Crittenden and N. Turok, *Phys. Rev. Lett.* **76**, 575 (1996); Planck Collab. 2015 Results XXI, *Astron. & Astrophys.* submitted, [arXiv:1502.01595](#).
41. W. Hu *et al.*, *Phys. Rev.* **D59**, 023512 (1999).
42. P.J.E. Peebles and J.T. Yu, *Astrophys. J.* **162**, 815 (1970); R.A. Sunyaev and Ya.B. Zeldovich, *Astrophys. & Space Sci.* **7**, 3 (1970).
43. D. Scott, J. Silk, and M. White, *Science* **268**, 829 (1995).
44. D.J. Eisenstein, *New Astron. Rev.* **49**, 360 (2005); W.J. Percival *et al.*, *MNRAS* **381**, 1053 (2007).
45. J. Silk, *Astrophys. J.* **151**, 459 (1968).
46. M. Zaldarriaga and U. Seljak, *Phys. Rev.* **D58**, 023003 (1998); A. Lewis and A. Challinor, *Phys. Reports* **429**, 1 (2006).
47. Planck Collab. 2013 Result XVII, *Astron. & Astrophys.* **571**, A17 (2014).
48. Planck Collab. 2015 Results XV, *Astron. & Astrophys.* submitted, [arXiv:1502.01591](#).
49. M. Kaplinghat, L. Knox, and Y.-S. Song, *Phys. Rev. Lett.* **91**, 241301 (2003).
50. Planck Collab. 2013 Results XV, *Astron. & Astrophys.* **571**, A15 (2014).
51. Planck Collab. 2015 Results XI, *Astron. & Astrophys.* submitted, [arXiv:1507.07283](#).
52. S. Das *et al.*, *JCAP* **04**, 014 (2014).
53. R. Keisler *et al.*, *Astrophys. J.* **743**, 28 (2011); C.L. Reichardt *et al.*, *Astrophys. J.* **755**, 70 (2012); K.T. Story *et al.*, *Astrophys. J.* **779**, 86 (2013).
54. W. Hu and M. White, *New Astron.* **2**, 323 (1997).
55. W. Hu and M. White, *Phys. Rev.* **D56**, 596 (1997).
56. M. Zaldarriaga and U. Seljak, *Phys. Rev.* **D55**, 1830 (1997); M. Kamionkowski, A. Kosowsky, and A. Stebbins, *Phys. Rev.* **D55**, 7368 (1997).
57. J. Kovac *et al.*, *Nature* **420**, 772 (2002).
58. D. Larson *et al.*, *Astrophys. J. Supp.* **192**, 16 (2011).
59. Keck Array V and BICEP2 Collabs, P.A.R. Ade *et al.*, *Astrophys. J.* **811**, 126, (2015).
60. S. Naess *et al.*, *JCAP* **10**, 007 (2014).
61. A.T. Crites *et al.*, *Astrophys. J.* **805**, 36 (2015).

62. E. Komatsu *et al.*, *Astrophys. J. Supp.* **192**, 18 (2011).
63. M.L. Brown *et al.*, *Astrophys. J.* **705**, 978 (2009).
64. D. Hanson *et al.*, *Phys. Rev. Lett.* **111**, 141301 (2013) [arXiv:1307.5830](#).
65. BICEP2 Collab., *Phys. Rev. Lett.* **112**, 1101 (2014).
66. BICEP/Keck and Planck Collabs., *Phys. Rev. Lett.* **114**, 1301 (2015).
67. POLARBEAR Collab., *Astrophys. J.* **794**, 171 (2014).
68. R. Keisler *et al.*, *Astrophys. J.* **807**, 151 (2015).
69. A. van Engelen *et al.*, *Astrophys. J.* **808**, 7 (2015).
70. Planck Collab. 2013 Results XII, *Astron. & Astrophys.* **571**, A12 (2014);  
Planck Collab. 2015 Results IX, *Astron. & Astrophys.* submit., [arXiv:1502.05956](#);  
Planck Collab. 2015 Results X, *Astron. & Astrophys.* submitted, [arXiv:1502.01588](#).
71. B. Gold *et al.*, *Astrophys. J. Supp.* **192**, 15 (2011).
72. Planck Collab. Interm. Results XXX, *Astron. & Astrophys.* submitted, [arXiv:1409.5738](#);  
Planck Collab. 2015 Results XII, *Astron. & Astrophys.* submitted, [arXiv:1506.06660](#).
73. M. Millea *et al.*, *Astrophys. J.* **746**, 4 (2012);  
E. Calabrese *et al.*, *JCAP* **8**, 010 (2014).
74. R.A. Sunyaev and Ya.B. Zeldovich, *Ann. Rev. Astron. Astrophys.* **18**, 537 (1980);  
M. Birkinshaw, *Phys. Reports* **310**, 98 (1999);  
J.E. Carlstrom, G.P. Holder, and E.D. Reese, *Ann. Rev. Astron. Astrophys.* **40**, 643 (2002).
75. R. Williamson *et al.*, *Astrophys. J.* **738**, 139 (2011);  
L.E. Bleem *et al.*, *Astrophys. J. Supp.* **216**, 27 (2015).
76. T.A. Marriage *et al.*, *Astrophys. J.* **737**, 61 (2011);  
M. Hasselfield *et al.*, *JCAP* **07**, 008 (2013).
77. Planck Collab. Early Results VIII, *Astron. & Astrophys.* **536**, 8 (2011);  
Planck Collab. 2013 Results XXIX, *Astron. & Astrophys.* **571**, A29 (2014);  
Planck Collab. 2015 Results XXVII, *Astron. & Astrophys.* submitted, [arXiv:1502.01598](#).
78. Planck Collab. 2013 Results XX, *Astron. & Astrophys.* **571**, A20 (2014);  
Planck Collab. 2015 Results XXIV, *Astron. & Astrophys.* submitted, [arXiv:1502.01597](#).
79. Planck Collab. 2013 Results XXIV, *Astron. & Astrophys.* **571**, A24 (2014);  
Planck Collab. 2015 Results XVII, *Astron. & Astrophys.* submitted, [arXiv:1502.01592](#).
80. F. Beutler *et al.*, *MNRAS* **416**, 3017 (2011);  
L. Anderson *et al.*, *MNRAS* **441**, 24 (2014);  
E.A. Kazin *et al.*, *MNRAS* **441**, 3524 (2014);  
A.J. Ross *et al.*, *MNRAS* **449**, 835 (2014).
81. Planck Collab. 2015 Results XIV, *Astron. & Astrophys.* submitted, [arXiv:1502.01590](#).
82. X. Fan, C.L. Carilli, and B. Keating, *Ann. Rev. Astron. Astrophys.* **44**, 415 (2006).
83. T.R. Choudhury *et al.*, *MNRAS* **452**, 261 (2015).
84. M. Kamionkowski and A. Kosowsky, *Ann. Rev. Nucl. and Part. Sci.* **49**, 77 (1999);  
A. Lasenby, *Space Sci. Rev.* **148**, 329 (2009).
85. Planck Collab. 2015 Results XIX, *Astron. & Astrophys.* submitted, [arXiv:1502.01594](#).
86. Planck Collab. Interm. Results XXIV, *Astron. & Astrophys.* **580**, A22 (2015).
87. R. Maartens, *Living Rev. Rel.* **7**, 7 (2004).

## 29. COSMIC RAYS

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### 29.1. Primary spectra

The cosmic radiation incident at the top of the terrestrial atmosphere includes all stable charged particles and nuclei with lifetimes of order  $10^6$  years or longer. Technically, “primary” cosmic rays are those particles accelerated at astrophysical sources and “secondaries” are those particles produced in interaction of the primaries with interstellar gas. Thus electrons, protons and helium, as well as carbon, oxygen, iron, and other nuclei synthesized in stars, are primaries. Nuclei such as lithium, beryllium, and boron (which are not abundant end-products of stellar nucleosynthesis) are secondaries. Antiprotons and positrons are also in large part secondary. Whether a small fraction of these particles may be primary is a question of current interest.

Apart from particles associated with solar flares, the cosmic radiation comes from outside the solar system. The incoming charged particles are “modulated” by the solar wind, the expanding magnetized plasma generated by the Sun, which decelerates and partially excludes the lower energy galactic cosmic rays from the inner solar system. There is a significant anticorrelation between solar activity (which has an alternating eleven-year cycle) and the intensity of the cosmic rays with energies below about 10 GeV. In addition, the lower-energy cosmic rays are affected by the geomagnetic field, which they must penetrate to reach the top of the atmosphere. Thus the intensity of any component of the cosmic radiation in the GeV range depends both on the location and time.

There are four different ways to describe the spectra of the components of the cosmic radiation: (1) By particles per unit rigidity. Propagation (and probably also acceleration) through cosmic magnetic fields depends on gyroradius or *magnetic rigidity*,  $R$ , which is gyroradius multiplied by the magnetic field strength:

$$R = \frac{pc}{Ze} = r_L B. \quad (29.1)$$

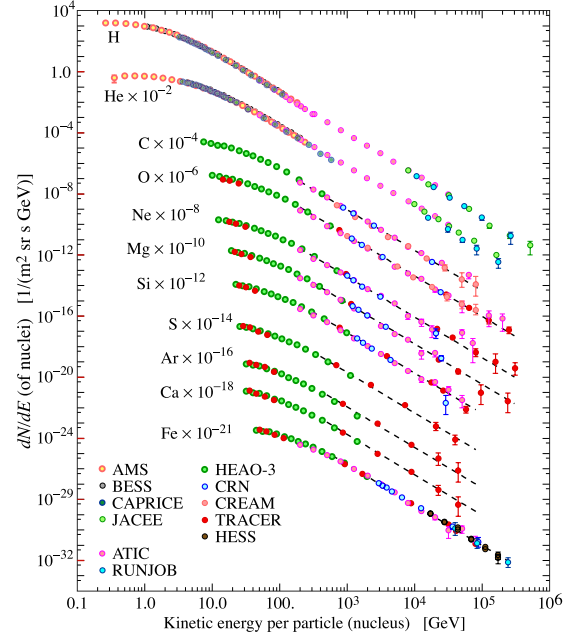
(2) By particles per energy-per-nucleon. Fragmentation of nuclei propagating through the interstellar gas depends on energy per nucleon, since that quantity is approximately conserved when a nucleus breaks up on interaction with the gas. (3) By nucleons per energy-per-nucleon. Production of secondary cosmic rays in the atmosphere depends on the intensity of nucleons per energy-per-nucleon, approximately independently of whether the incident nucleons are free protons or bound in nuclei. (4) By particles per energy-per-nucleus. Air shower experiments that use the atmosphere as a calorimeter generally measure a quantity that is related to total energy per particle.

The units of differential intensity  $I$  are  $[\text{m}^{-2} \text{s}^{-1} \text{sr}^{-1} \mathcal{E}^{-1}]$ , where  $\mathcal{E}$  represents the units of one of the four variables listed above.

The intensity of primary nucleons in the energy range from several GeV to somewhat beyond 100 TeV is given approximately by

$$I_N(E) \approx 1.8 \times 10^4 (E/1 \text{ GeV})^{-\alpha} \frac{\text{nucleons}}{\text{m}^2 \text{ s sr GeV}}, \quad (29.2)$$

where  $E$  is the energy-per-nucleon (including rest mass energy) and  $\alpha$  ( $\equiv \gamma + 1$ ) = 2.7 is the differential spectral index of the cosmic-ray flux and  $\gamma$  is the integral spectral index. About 79% of the primary nucleons are free protons and about 70% of the rest are nucleons bound in helium nuclei. The fractions of the primary nuclei are nearly constant over this energy range (possibly with small but interesting variations). Fractions of both primary and secondary incident nuclei are listed in Table 29.1. Figure 29.1 shows the major components for energies greater than 2 GeV/nucleon. A useful compendium of experimental data for cosmic-ray nuclei and electrons is described in [1].



**Figure 29.1:** Fluxes of nuclei of the primary cosmic radiation in particles per energy-per-nucleus are plotted vs energy-per-nucleus using data from Refs. [2–13]. The figure was created by P. Boyle and D. Muller.

The composition and energy spectra of nuclei are typically interpreted in the context of propagation models, in which the sources of the primary cosmic radiation are located within the Galaxy [14]. The ratio of secondary to primary nuclei is observed to decrease with increasing energy, a fact interpreted to mean that the lifetime of cosmic rays in the galaxy decreases with energy. Measurements of radioactive “clock” isotopes in the low energy cosmic radiation are consistent with a lifetime in the galaxy of about 15 Myr [15].

**Table 29.1:** Relative abundances  $F$  of cosmic-ray nuclei at 10.6 GeV/nucleon normalized to oxygen ( $\equiv 1$ ) [7]. The oxygen flux at kinetic energy of 10.6 GeV/nucleon is  $3.29 \times 10^{-2} (\text{m}^2 \text{s sr GeV/nucleon})^{-1}$ . Abundances of hydrogen and helium are from Refs. [3,4]. Note that one can not use these values to extend the cosmic-ray flux to high energy because the power law indices for each element may differ slightly.

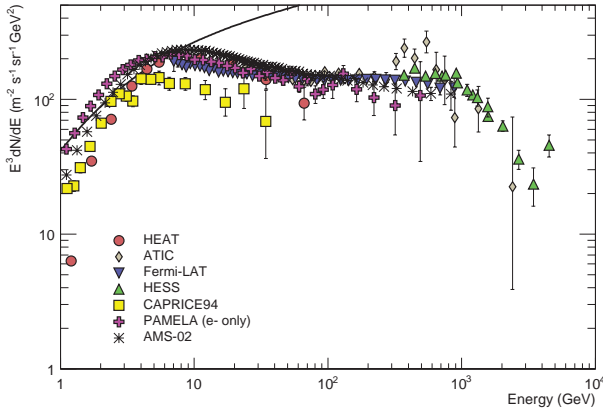
$Z$	Element	$F$	$Z$	Element	$F$
1	H	540	13–14	Al-Si	0.19
2	He	26	15–16	P-S	0.03
3–5	Li-B	0.40	17–18	Cl-Ar	0.01
6–8	C-O	2.20	19–20	K-Ca	0.02
9–10	F-Ne	0.30	21–25	Sc-Mn	0.05
11–12	Na-Mg	0.22	26–28	Fe-Ni	0.12

Cosmic rays are nearly isotropic at most energies due to diffusive propagation in the galactic magnetic field. Milagro [16], IceCube [17], and the Tibet-III air shower array [18] have observed anisotropy at the level of about  $10^{-3}$  for cosmic rays with energy of a few TeV, possibly due to nearby sources.

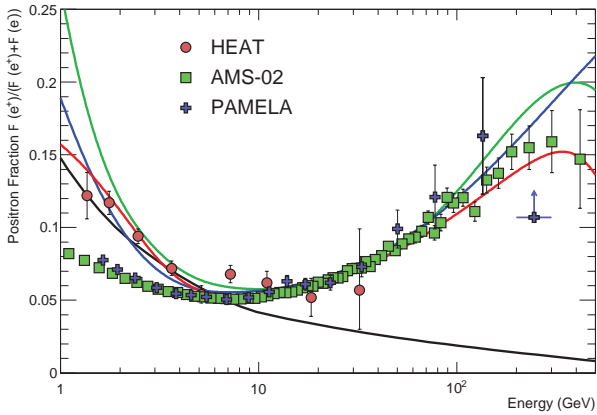
The spectrum of electrons and positrons incident at the top of the atmosphere is expected to steepen by one power of  $E$  at an energy of  $\sim 5$  GeV because of strong radiative energy loss effects in the galaxy. The ATIC experiment [19] measured a sharp excess of electrons over propagation model expectations, at energies of  $\sim 300$ –800 GeV. The *Fermi*/LAT  $\gamma$ -ray observatory measured a not-entirely flat

spectrum [20] without confirming the peak of the ATIC excess at  $\sim 600$  GeV. Measurements in the same energy range by AMS-02 also show no sharp features and are compatible with a single power law above 30.2 GeV [21]. The HESS imaging atmospheric Cherenkov array also measured the electron flux above  $\sim 400$  GeV, finding indications of a cutoff above  $\sim 1$  TeV [22], but no evidence for a pronounced peak below this.

The PAMELA [26] and AMS-02 [27,24] satellite experiments measured the positron to electron ratio to increase above 10 GeV instead of the expected decrease [28] at higher energy, confirming earlier hints seen by the HEAT balloon-borne experiment [30]. The structure in the electron spectrum, as well as the increase in the positron fraction, may be related to contributions from individual nearby sources (supernova remnants or pulsars) emerging above a background suppressed at high energy by synchrotron losses [31]. Other explanations have invoked propagation effects [32] or dark matter decay/annihilation processes (see, e.g., [29]). The significant disagreement in the ratio below  $\sim 10$  GeV is attributable to differences in charge-sign dependent solar modulation effects present near earth at the times of measurement.



**Figure 29.2:** Differential spectrum of electrons plus positrons (except PAMELA data, which are electrons only) multiplied by  $E^3$  [19–23,33,34]. The line shows the proton spectrum [25] multiplied by 0.01.



**Figure 29.3:** The positron fraction (ratio of the flux of  $e^+$  to the total flux of  $e^+$  and  $e^-$ ) [26,24,30]. The heavy black line is a model of pure secondary production [28] and the three thin lines show three representative attempts to model the positron excess with different phenomena: green: dark matter decay [29]; blue: propagation physics [32]; red: production in pulsars [40]. The ratio below 10 GeV is dependent on the polarity of the solar magnetic field.

The ratio of antiprotons to protons is  $\sim 2 \times 10^{-4}$  [35] at around 10–20 GeV, and there is clear evidence [36] for the kinematic suppression at lower energy that is the signature of secondary antiprotons. The  $\bar{p}/p$  ratio also shows a strong dependence on the phase and polarity of the solar cycle [37] in the opposite sense to that of the positron fraction. There is at this time no evidence for a significant primary component of antiprotons. No antihelium or antideuteron has been found in the cosmic radiation. The best measured upper limit on the ratio antihelium/helium is currently approximately  $1 \times 10^{-7}$  [38]. The upper limit on the flux of antideuterons around 1 GeV/nucleon is approximately  $2 \times 10^{-4} (\text{m}^2 \text{ s sr GeV/nucleon})^{-1}$  [39].

## 29.2. Cosmic rays in the atmosphere

Figure 29.4 shows the vertical fluxes of the major cosmic-ray components in the atmosphere in the energy region where the particles are most numerous (except for electrons, which are most numerous near their critical energy, which is about 81 MeV in air). Except for protons and electrons near the top of the atmosphere, all particles are produced in interactions of the primary cosmic rays in the air. Muons and neutrinos are products of the decay chain of charged mesons, while electrons and photons originate in decays of neutral mesons.

Most measurements are made at ground level or near the top of the atmosphere, but there are also measurements of muons and electrons from airplanes and balloons. Fig. 29.4 includes recent measurements of negative muons [41–45]. Since  $\mu^+(\mu^-)$  are produced in association with  $\nu_\mu(\bar{\nu}_\mu)$ , the measurement of muons near the maximum of the intensity curve for the parent pions serves to calibrate the atmospheric  $\nu_\mu$  beam [46]. Because muons typically lose almost 2 GeV in passing through the atmosphere, the comparison near the production altitude is important for the sub-GeV range of  $\nu_\mu(\bar{\nu}_\mu)$  energies.

The flux of cosmic rays through the atmosphere is described by a set of coupled cascade equations with boundary conditions at the top of the atmosphere to match the primary spectrum. Numerical or Monte Carlo calculations are needed to account accurately for decay and energy-loss processes, and for the energy-dependences of the cross sections and of the primary spectral index  $\gamma$ . Approximate analytic solutions are, however, useful in limited regions of energy [47,48]. For example, the vertical intensity of charged pions with energy  $E_\pi \ll \epsilon_\pi = 115$  GeV is

$$I_\pi(E_\pi, X) \approx \frac{Z_{N\pi}}{\lambda_N} I_N(E_\pi, 0) e^{-X/\Lambda} \frac{X E_\pi}{\epsilon_\pi}, \quad (29.3)$$

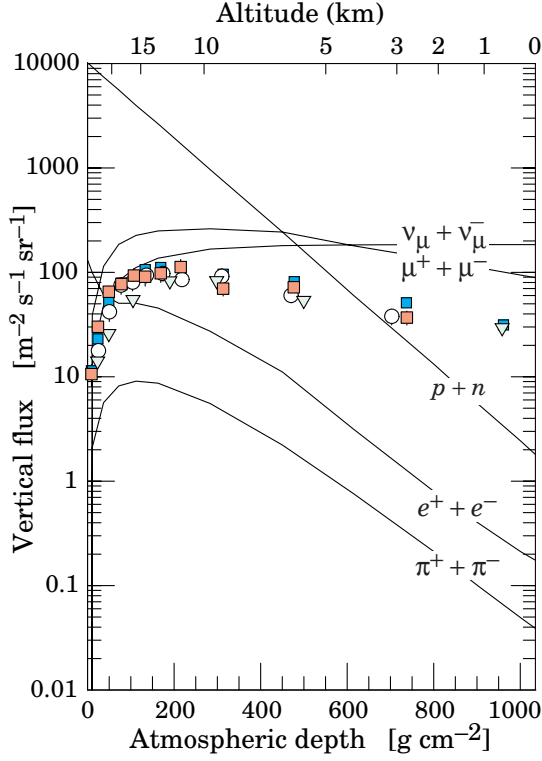
where  $\Lambda$  is the characteristic length for exponential attenuation of the parent nucleon flux in the atmosphere. This expression has a maximum at  $X = \Lambda \approx 121 \pm 4 \text{ g cm}^{-2}$  [49], which corresponds to an altitude of 15 kilometers. The quantity  $Z_{N\pi}$  is the spectrum-weighted moment of the inclusive distribution of charged pions in interactions of nucleons with nuclei of the atmosphere. The intensity of low-energy pions is much less than that of nucleons because  $Z_{N\pi} \approx 0.079$  is small and because most pions with energy much less than the critical energy  $\epsilon_\pi$  decay rather than interact.

## 29.3. Cosmic rays at the surface

**29.3.1. Muons:** Muons are the most numerous charged particles at sea level (see Fig. 29.4). Most muons are produced high in the atmosphere (typically 15 km) and lose about 2 GeV to ionization before reaching the ground. Their energy and angular distribution reflect a convolution of the production spectrum, energy loss in the atmosphere, and decay. For example, 2.4 GeV muons have a decay length of 15 km, which is reduced to 8.7 km by energy loss. The mean energy of muons at the ground is  $\approx 4$  GeV. The energy spectrum is almost flat below 1 GeV, steepens gradually to reflect the primary spectrum in the 10–100 GeV range, and steepens further at higher energies because pions with  $E_\pi > \epsilon_\pi$  tend to interact in the atmosphere before they decay. Asymptotically ( $E_\mu \gg 1$  TeV), the energy spectrum of atmospheric muons is one power steeper than the primary spectrum. The integral intensity of vertical muons above 1 GeV/c at sea level is  $\approx 70 \text{ m}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$  [50,51], with recent measurements [52–54] favoring a lower normalization by 10–15%. Experimentalists are familiar with



this number in the form  $I \approx 1 \text{ cm}^{-2} \text{ min}^{-1}$  for horizontal detectors. The overall angular distribution of muons at the ground is  $\propto \cos^2 \theta$ , which is characteristic of muons with  $E_\mu \sim 3 \text{ GeV}$ . At lower energy the angular distribution becomes increasingly steep, while at higher energy it flattens, approaching a  $\sec \theta$  distribution for  $E_\mu \gg \epsilon_\pi$  and  $\theta < 70^\circ$ .



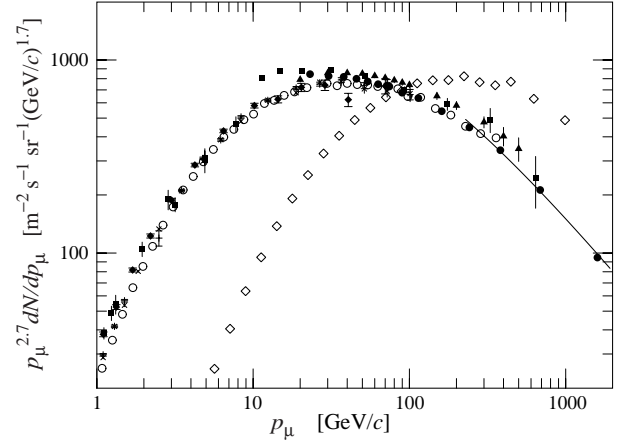
**Figure 29.4:** Vertical fluxes of cosmic rays in the atmosphere with  $E > 1 \text{ GeV}$  estimated from the nucleon flux of Eq. (29.2). The points show measurements of negative muons with  $E_\mu > 1 \text{ GeV}$  [41–45].

Figure 29.5 shows the muon energy spectrum at sea level for two angles. At large angles low energy muons decay before reaching the surface and high energy pions decay before they interact, thus the average muon energy increases. An approximate extrapolation formula valid when muon decay is negligible ( $E_\mu > 100/\cos \theta \text{ GeV}$ ) and the curvature of the Earth can be neglected ( $\theta < 70^\circ$ ) is

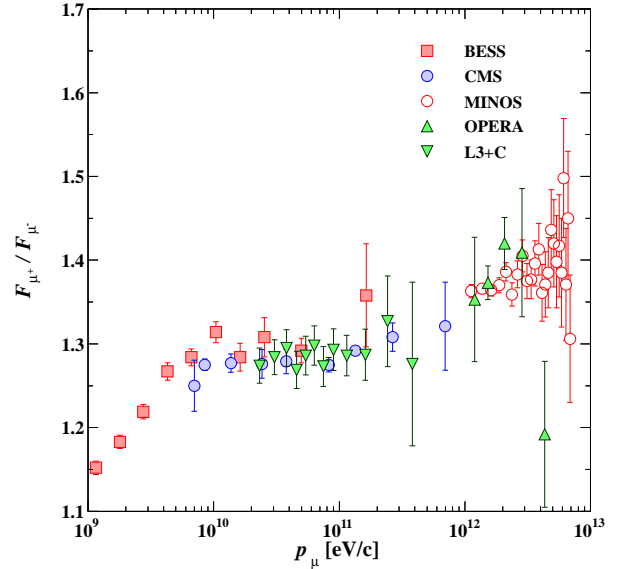
$$\frac{dN_\mu}{dE_\mu d\Omega} \approx \frac{0.14 E_\mu^{-2.7}}{\text{cm}^2 \text{ s sr GeV}} \times \left\{ \frac{1}{1 + \frac{1.1 E_\mu \cos \theta}{115 \text{ GeV}}} + \frac{0.054}{1 + \frac{1.1 E_\mu \cos \theta}{850 \text{ GeV}}} \right\}, \quad (29.4)$$

where the two terms give the contribution of pions and charged kaons. Eq. (29.4) neglects a small contribution from charm and heavier flavors which is negligible except at very high energy [55].

The muon charge ratio reflects the excess of  $\pi^+$  over  $\pi^-$  and  $K^+$  over  $K^-$  in the forward fragmentation region of proton initiated interactions together with the fact that there are more protons than neutrons in the primary spectrum. The increase with energy of  $\mu^+/\mu^-$  shown in Fig. 29.6 reflects the increasing importance of kaons in the TeV range [60] and indicates a significant contribution of associated production by cosmic-ray protons ( $p \rightarrow \Lambda + K^+$ ). The same process is even more important for atmospheric neutrinos at high energy.



**Figure 29.5:** Spectrum of muons at  $\theta = 0^\circ$  ( $\diamond$  [50],  $\blacksquare$  [56],  $\blacktriangledown$  [57],  $\blacktriangle$  [58],  $\times$ ,  $+$  [52],  $\circ$  [53], and  $\bullet$  [54] and  $\theta = 75^\circ$   $\diamond$  [59]). The line plots the result from Eq. (29.4) for vertical showers.



**Figure 29.6:** Muon charge ratio as a function of the muon momentum from Refs. [53,54,60,65,66].

**29.3.2. Electromagnetic component:** At the ground, this component consists of electrons, positrons, and photons primarily from cascades initiated by decay of neutral and charged mesons. Muon decay is the dominant source of low-energy electrons at sea level. Decay of neutral pions is more important at high altitude or when the energy threshold is high. Knock-on electrons also make a small contribution at low energy [61]. The integral vertical intensity of electrons plus positrons is very approximately 30, 6, and  $0.2 \text{ m}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$  above 10, 100, and 1000 MeV respectively [51,62], but the exact numbers depend sensitively on altitude, and the angular dependence is complex because of the different altitude dependence of the different sources of electrons [61–63]. The ratio of photons to electrons plus positrons is approximately 1.3 above 1 GeV and 1.7 below the critical energy [63].

**29.3.3. Protons :** Nucleons above 1 GeV/c at ground level are degraded remnants of the primary cosmic radiation. The intensity is approximately  $I_N(E, 0) \times \exp(-X/\cos\theta\Lambda)$  for  $\theta < 70^\circ$ . At sea level, about 1/3 of the nucleons in the vertical direction are neutrons (up from  $\approx 10\%$  at the top of the atmosphere as the  $n/p$  ratio approaches equilibrium). The integral intensity of vertical protons above 1 GeV/c at sea level is  $\approx 0.9 \text{ m}^{-2}\text{s}^{-1}\text{sr}^{-1}$  [51,64].

## 29.4. Cosmic rays underground

Only muons and neutrinos penetrate to significant depths underground. The muons produce tertiary fluxes of photons, electrons, and hadrons.

**29.4.1. Muons :** As discussed in Section 33.6 of this *Review*, muons lose energy by ionization and by radiative processes: bremsstrahlung, direct production of  $e^+e^-$  pairs, and photonuclear interactions. The total muon energy loss may be expressed as a function of the amount of matter traversed as

$$-\frac{dE_\mu}{dX} = a + b E_\mu, \quad (29.5)$$

where  $a$  is the ionization loss and  $b$  is the fractional energy loss by the three radiation processes. Both are slowly varying functions of energy. The quantity  $\epsilon \equiv a/b$  ( $\approx 500$  GeV in standard rock) defines a critical energy below which continuous ionization loss is more important than radiative losses. Table 29.2 shows  $a$  and  $b$  values for standard rock, and  $b$  for ice, as a function of muon energy. The second column of Table 29.2 shows the muon range in standard rock ( $A = 22$ ,  $Z = 11$ ,  $\rho = 2.65 \text{ g cm}^{-3}$ ). These parameters are quite sensitive to the chemical composition of the rock, which must be evaluated for each location.

**Table 29.2:** Average muon range  $R$  and energy loss parameters  $a$  and  $b$  calculated for standard rock [67] and the total energy loss parameter  $b$  for ice. Range is given in km-water-equivalent, or  $10^5 \text{ g cm}^{-2}$ .

$E_\mu$ GeV	$R$ km.w.e.	$a$ MeV $\text{g}^{-1} \text{cm}^2$	$b_{\text{brems}}$ —	$b_{\text{pair}}$ $10^{-6} \text{ g}^{-1} \text{cm}^2$	$b_{\text{nucl}}$ —	$\sum b_i$ —	$\sum b(\text{ice})$ —
10	0.05	2.17	0.70	0.70	0.50	1.90	1.66
100	0.41	2.44	1.10	1.53	0.41	3.04	2.51
1000	2.45	2.68	1.44	2.07	0.41	3.92	3.17
10000	6.09	2.93	1.62	2.27	0.46	4.35	3.78

The intensity of muons underground can be estimated from the muon intensity in the atmosphere and their rate of energy loss. To the extent that the mild energy dependence of  $a$  and  $b$  can be neglected, Eq. (29.5) can be integrated to provide the following relation between the energy  $E_{\mu,0}$  of a muon at production in the atmosphere and its average energy  $E_\mu$  after traversing a thickness  $X$  of rock (or ice or water):

$$E_{\mu,0} = (E_\mu + \epsilon) e^{bX} - \epsilon. \quad (29.6)$$

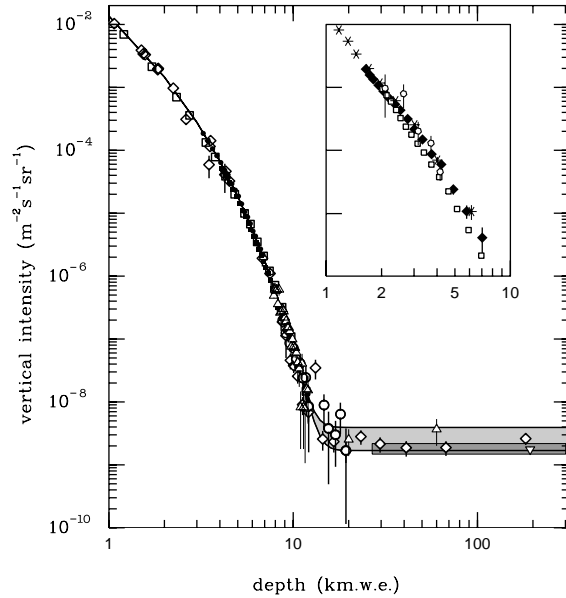
Especially at high energy, however, fluctuations are important and an accurate calculation requires a simulation that accounts for stochastic energy-loss processes [68].

There are two depth regimes for which Eq. (29.6) can be simplified. For  $X \ll b^{-1} \approx 2.5$  km water equivalent,  $E_{\mu,0} \approx E_\mu(X) + aX$ , while for  $X \gg b^{-1}$   $E_{\mu,0} \approx (\epsilon + E_\mu(X)) \exp(bX)$ . Thus at shallow depths the differential muon energy spectrum is approximately constant for  $E_\mu < aX$  and steepens to reflect the surface muon spectrum for  $E_\mu > aX$ , whereas for  $X > 2.5$  km.w.e. the differential spectrum underground is again constant for small muon energies but steepens to reflect the surface muon spectrum for  $E_\mu > \epsilon \approx 0.5$  TeV. In the deep regime the shape is independent of depth although the intensity decreases exponentially with depth. In general the muon spectrum at slant depth  $X$  is

$$\frac{dN_\mu(X)}{dE_\mu} = \frac{dN_\mu}{dE_{\mu,0}} \frac{dE_{\mu,0}}{dE_\mu} = \frac{dN_\mu}{dE_{\mu,0}} e^{bX}, \quad (29.7)$$

where  $E_{\mu,0}$  is the solution of Eq. (29.6) in the approximation neglecting fluctuations.

Fig. 29.7 shows the vertical muon intensity versus depth. In constructing this “depth-intensity curve,” each group has taken account of the angular distribution of the muons in the atmosphere, the map of the overburden at each detector, and the properties of the local medium in connecting measurements at various slant depths and zenith angles to the vertical intensity. Use of data from a range of angles allows a fixed detector to cover a wide range of depths. The flat portion of the curve is due to muons produced locally by charged-current interactions of  $\nu_\mu$ . The inset shows the vertical intensity curve for water and ice published in Refs. [70–73]. It is not as steep as the one for rock because of the lower muon energy loss in water.



**Figure 29.7:** Vertical muon intensity vs depth (1 km.w.e. =  $10^5 \text{ g cm}^{-2}$  of standard rock). The experimental data are from:  $\diamond$ : the compilations of Crouch [69],  $\square$ : Baksan [74],  $\circ$ : LVD [75],  $\bullet$ : MACRO [76],  $\blacksquare$ : Frejus [77], and  $\triangle$ : SNO [78]. The shaded area at large depths represents neutrino-induced muons of energy above 2 GeV. The upper line is for horizontal neutrino-induced muons, the lower one for vertically upward muons. Darker shading shows the muon flux measured by the SuperKamiokande experiment. The inset shows the vertical intensity curve for water and ice published in Refs. [70–73].

### 29.4.2. Neutrinos :

Because neutrinos have small interaction cross sections, measurements of atmospheric neutrinos require a deep detector to avoid backgrounds. There are two types of measurements: contained (or semi-contained) events, in which the vertex is determined to originate inside the detector, and neutrino-induced muons. The latter are muons that enter the detector from zenith angles so large (e.g., nearly horizontal or upward) that they cannot be muons produced in the atmosphere. In neither case is the neutrino flux measured directly. What is measured is a convolution of the neutrino flux and cross section with the properties of the detector (which includes the surrounding medium in the case of entering muons).

Contained and semi-contained events reflect neutrinos in the sub-GeV to multi-GeV region where the product of increasing cross section and decreasing flux is maximum. In the GeV region the neutrino flux and its angular distribution depend on the geomagnetic location of the detector and, to a lesser extent, on the phase of the solar cycle. Naively, we expect  $\nu_\mu/\nu_e = 2$  from counting neutrinos of the two flavors coming from the chain of pion and muon decay. Contrary to expectation, however, the numbers of the two classes of



**Table 29.3:** Measured fluxes ( $10^{-9} \text{ m}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$ ) of neutrino-induced muons as a function of the effective minimum muon energy  $E_\mu$ .

$E_\mu >$	1 GeV	1 GeV	1 GeV	2 GeV	3 GeV	3 GeV
Ref.	CWI [82]	Baksan [83]	MACRO [84]	IMB [85]	Kam [86]	SuperK [87]
$F_\mu$	$2.17 \pm 0.21$	$2.77 \pm 0.17$	$2.29 \pm 0.15$	$2.26 \pm 0.11$	$1.94 \pm 0.12$	$1.74 \pm 0.07$

events are similar rather than different by a factor of two. This is now understood to be a consequence of neutrino flavor oscillations [81]. (See the article on neutrino properties in this *Review*.)

Two well-understood properties of atmospheric cosmic rays provide a standard for comparison of the measurements of atmospheric neutrinos to expectation. These are the “sec  $\theta$  effect” and the “east-west effect” [80]. The former refers originally to the enhancement of the flux of  $> 10$  GeV muons (and neutrinos) at large zenith angles because the parent pions propagate more in the low density upper atmosphere where decay is enhanced relative to interaction. For neutrinos from muon decay, the enhancement near the horizontal becomes important for  $E_\nu > 1$  GeV and arises mainly from the increased pathlength through the atmosphere for muon decay in flight. Fig. 14.11 from Ref. 79 shows a comparison between measurement and expectation for the zenith angle dependence of multi-GeV electron-like (mostly  $\nu_e$ ) and muon-like (mostly  $\nu_\mu$ ) events separately. The  $\nu_e$  show an enhancement near the horizontal and approximate equality for nearly upward ( $\cos \theta \approx -1$ ) and nearly downward ( $\cos \theta \approx 1$ ) events. There is, however, a very significant deficit of upward ( $\cos \theta < 0$ )  $\nu_\mu$  events, which have long pathlengths comparable to the radius of the Earth. This feature is the principal signature for atmospheric neutrino oscillations [81].

Muons that enter the detector from outside after production in charged-current interactions of neutrinos naturally reflect a higher energy portion of the neutrino spectrum than contained events because the muon range increases with energy as well as the cross section. The relevant energy range is  $\sim 10 < E_\nu < 1000$  GeV, depending somewhat on angle. Neutrinos in this energy range show a sec  $\theta$  effect similar to muons (see Eq. (29.4)). This causes the flux of horizontal neutrino-induced muons to be approximately a factor two higher than the vertically upward flux. The upper and lower edges of the horizontal shaded region in Fig. 29.7 correspond to horizontal and vertical intensities of neutrino-induced muons. Table 29.3 gives the measured fluxes of upward-moving neutrino-induced muons averaged over the lower hemisphere. Generally the definition of minimum muon energy depends on where it passes through the detector. The tabulated effective minimum energy estimates the average over various accepted trajectories.

## 29.5. Air showers

So far we have discussed inclusive or uncorrelated fluxes of various components of the cosmic radiation. An air shower is caused by a single cosmic ray with energy high enough for its cascade to be detectable at the ground. The shower has a hadronic core, which acts as a collimated source of electromagnetic subshowers, generated mostly from  $\pi^0 \rightarrow \gamma\gamma$  decays. The resulting electrons and positrons are the most numerous charged particles in the shower. The number of muons, produced by decays of charged mesons, is an order of magnitude lower. Air showers spread over a large area on the ground, and arrays of detectors operated for long times are useful for studying cosmic rays with primary energy  $E_0 > 100$  TeV, where the low flux makes measurements with small detectors in balloons and satellites difficult.

Greisen [88] gives the following approximate expressions for the numbers and lateral distributions of particles in showers at ground level. The total number of muons  $N_\mu$  with energies above 1 GeV is

$$N_\mu(> 1 \text{ GeV}) \approx 0.95 \times 10^5 \left( N_e / 10^6 \right)^{3/4}, \quad (29.8)$$

where  $N_e$  is the total number of charged particles in the shower (not just  $e^\pm$ ). The number of muons per square meter,  $\rho_\mu$ , as a function of the lateral distance  $r$  (in meters) from the center of the shower is

$$\rho_\mu = \frac{1.25 N_\mu}{2\pi \Gamma(1.25)} \left( \frac{1}{320} \right)^{1.25} r^{-0.75} \left( 1 + \frac{r}{320} \right)^{-2.5}, \quad (29.9)$$

where  $\Gamma$  is the gamma function. The number density of charged particles is

$$\rho_e = C_1(s, d, C_2) x^{(s-2)} (1+x)^{(s-4.5)} (1+C_2 x^d). \quad (29.10)$$

Here  $s$ ,  $d$ , and  $C_2$  are parameters in terms of which the overall normalization constant  $C_1(s, d, C_2)$  is given by

$$C_1(s, d, C_2) = \frac{N_e}{2\pi r_1^2} [B(s, 4.5 - 2s) + C_2 B(s + d, 4.5 - d - 2s)]^{-1}, \quad (29.11)$$

where  $B(m, n)$  is the beta function. The values of the parameters depend on shower size ( $N_e$ ), depth in the atmosphere, identity of the primary nucleus, etc. For showers with  $N_e \approx 10^6$  at sea level, Greisen uses  $s = 1.25$ ,  $d = 1$ , and  $C_2 = 0.088$ . Finally,  $x = r/r_1$ , where  $r_1$  is the Molière radius, which depends on the density of the atmosphere and hence on the altitude at which showers are detected. At sea level  $r_1 \approx 78$  m. It increases with altitude as the air density decreases. (See the section on electromagnetic cascades in the article on the passage of particles through matter in this *Review*.)

The lateral spread of a shower is determined largely by Coulomb scattering of the many low-energy electrons and is characterized by the Molière radius. The lateral spread of the muons ( $\rho_\mu$ ) is larger and depends on the transverse momenta of the muons at production as well as multiple scattering.

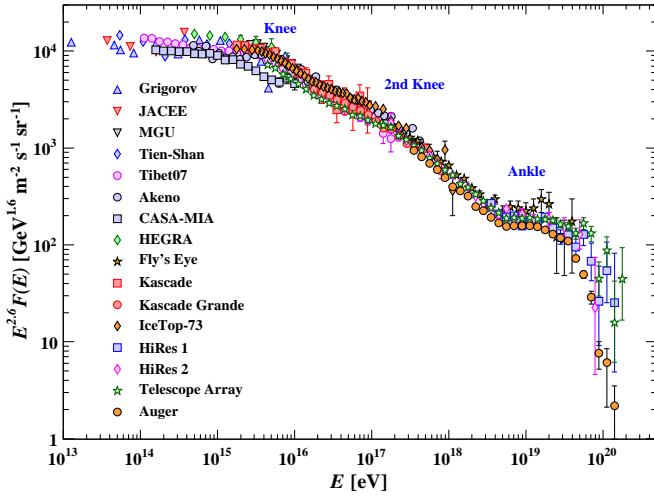
There are large fluctuations in development from shower to shower, even for showers of the same energy and primary mass—especially for small showers, which are usually well past maximum development when observed at the ground. Thus the shower size  $N_e$  and primary energy  $E_0$  are only related in an average sense, and even this relation depends on depth in the atmosphere. One estimate of the relation is [95]

$$E_0 \sim 3.9 \times 10^6 \text{ GeV} (N_e / 10^6)^{0.9} \quad (29.12)$$

for vertical showers with  $10^{14} < E < 10^{17}$  eV at 920 g cm $^{-2}$  (965 m above sea level). As  $E_0$  increases the shower maximum (on average) moves down into the atmosphere and the relation between  $N_e$  and  $E_0$  changes. Moreover, because of fluctuations,  $N_e$  as a function of  $E_0$  is not correctly obtained by inverting Eq. (29.12). At the maximum of shower development, there are approximately 2/3 particles per GeV of primary energy.

There are three common types of air shower detectors: shower arrays that study the shower size  $N_e$  and the lateral distribution on the ground, Cherenkov detectors that detect the Cherenkov radiation emitted by the charged particles of the shower, and fluorescence detectors that study the nitrogen fluorescence excited by the charged particles in the shower. The fluorescence light is emitted isotropically so the showers can be observed from the side. Detailed simulations and cross-calibrations between different types of detectors are necessary to establish the primary energy spectrum from air-shower experiments.

Figure 29.8 shows the “all-particle” spectrum. The differential energy spectrum has been multiplied by  $E^{2.6}$  in order to display the features of the steep spectrum that are otherwise difficult to discern. The steepening that occurs between  $10^{15}$  and  $10^{16}$  eV is known as the *knee* of the spectrum. The feature around  $10^{18.5}$  eV is called the *ankle* of the spectrum.



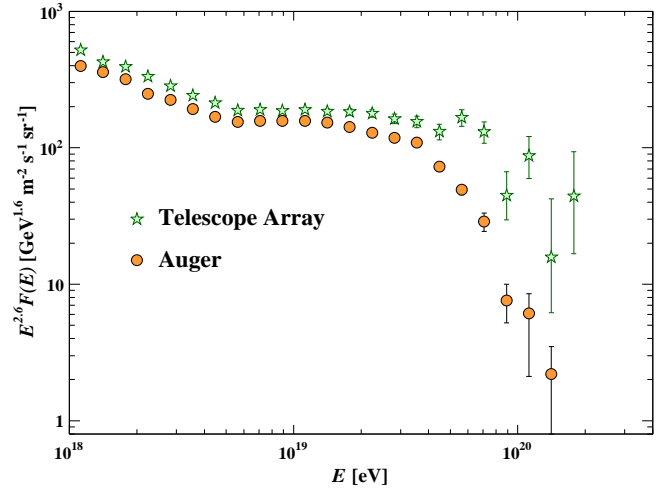
**Figure 29.8:** The all-particle spectrum as a function of  $E$  (energy-per-nucleus) from air shower measurements [90–105].

Measurements of flux with small air shower experiments in the knee region differ by as much as a factor of two, indicative of systematic uncertainties in interpretation of the data. (For a review see Ref. 89.) In establishing the spectrum shown in Fig. 29.8, efforts have been made to minimize the dependence of the analysis on the primary composition. Ref. 98 uses an unfolding procedure to obtain the spectra of the individual components, giving a result for the all-particle spectrum between  $10^{15}$  and  $10^{17}$  eV that lies toward the upper range of the data shown in Fig. 29.8. In the energy range above  $10^{17}$  eV, the fluorescence technique [106] is particularly useful because it can establish the primary energy in a model-independent way by observing most of the longitudinal development of each shower, from which  $E_0$  is obtained by integrating the energy deposition in the atmosphere. The result, however, depends strongly on the light absorption in the atmosphere and the calculation of the detector's aperture.

Assuming the cosmic-ray spectrum below  $10^{18}$  eV is of galactic origin, the *knee* could reflect the fact that most cosmic accelerators in the galaxy have reached their maximum energy. Some types of expanding supernova remnants, for example, are estimated not to be able to accelerate protons above energies in the range of  $10^{15}$  eV. Effects of propagation and confinement in the galaxy [109] also need to be considered. The Cascade-Grande experiment [100] has reported observation of a second steepening of the spectrum near  $8 \times 10^{16}$  eV, with evidence that this structure is accompanied a transition to heavy primaries.

Concerning the ankle, one possibility is that it is the result of a higher energy population of particles overtaking a lower energy population, for example an extragalactic flux beginning to dominate over the galactic flux (e.g. Ref. 106). Another possibility is that the dip structure in the region of the ankle is due to  $p\gamma \rightarrow e^+ + e^-$  energy losses of extragalactic protons on the 2.7 K cosmic microwave radiation (CMB) [111]. This dip structure has been cited as a robust signature of both the protonic and extragalactic nature of the highest energy cosmic rays [110]. If this interpretation is correct, then the galactic cosmic rays do not contribute significantly to the flux above  $10^{18}$  eV, consistent with the maximum expected range of acceleration by supernova remnants.

The energy-dependence of the composition from the knee through the ankle is useful in discriminating between these two viewpoints, since a heavy composition above  $10^{18}$  eV is inconsistent with the formation of the ankle by pair production losses on the CMB. The HiRes and Auger experiments, however, present very different interpretations of data on the depth of shower maximum  $X_{max}$ , a quantity that correlates strongly with the interaction cross section of the primary particle. If these results are interpreted using standard extrapolations of measured proton and nuclear cross sections, then the



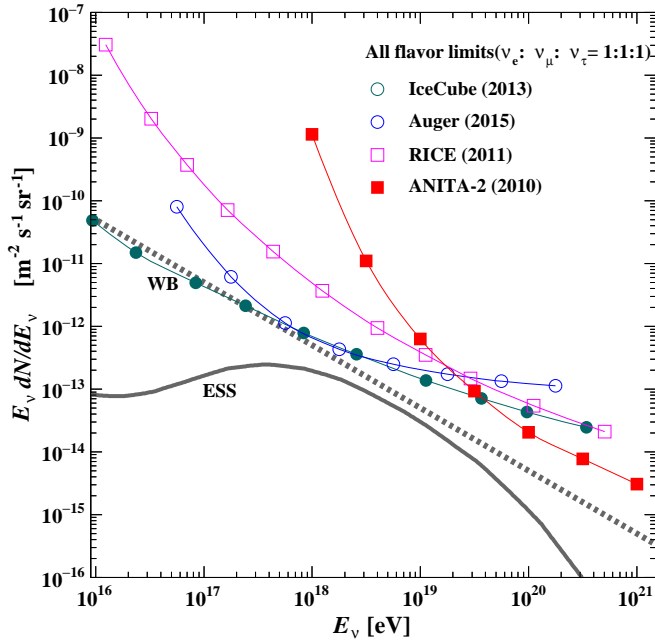
**Figure 29.9:** Expanded view of the highest energy portion of the cosmic-ray spectrum from data of the Telescope Array [104], and the Auger Observatory [105].

HiRes data [112] is consistent with the ultrahigh-energy cosmic-ray (UHECR) composition getting lighter and containing only protons and helium above  $10^{19}$  eV, while Auger [113,114] sees a composition getting lighter up to  $2 \times 10^{18}$  eV and becoming heavier after that, intermediate between protons and iron at  $3 \times 10^{19}$  eV. This may mean that the extragalactic cosmic rays have a mixed composition at acceleration similar to the GeV galactic cosmic rays. It is important to note that the measurements of  $X_{max}$  may be interpreted with equal validity in terms of a changing proton-air cross-section and no change in composition.

If the cosmic-ray flux at the highest energies is cosmological in origin, there should be a rapid steepening of the spectrum (called the GZK feature) around  $5 \times 10^{19}$  eV, resulting from the onset of inelastic interactions of UHE cosmic rays with the cosmic microwave background [115,116]. Photo-dissociation of heavy nuclei in the mixed composition model [117] would have a similar effect. UHECR experiments have detected events of energy above  $10^{20}$  eV [106–107]. The HiRes fluorescence experiment [102,125] detected evidence of the GZK suppression, and the Auger observatory [103–105] has also presented spectra showing this suppression based on surface detector measurements calibrated against fluorescence detectors using events detected in hybrid mode, i.e. with both the surface and the fluorescence detectors. The Telescope Array (TA) [104] has also presented a spectrum showing this suppression. The differential energy spectra measured by the TA and by Auger agree within systematic errors below  $10^{19}$  eV (Fig. 29.9). At higher energies, TA observes more events than would be expected if the spectral shape were the same as that seen by Auger. TA has also reported a ‘hot spot’ in the Northern Hemisphere at energies above  $5.5 \times 10^{19}$  eV of radius  $\sim 20^\circ$  with a post-trials statistical significance of this excess with respect to an isotropic distribution of  $3.4\sigma$  [108].

One half of the energy that UHECR protons lose in photoproduction interactions that cause the GZK effects ends up in neutrinos [118]. Measuring this *cosmogenic* neutrino flux above  $10^{18}$  eV would help resolve the UHECR uncertainties mentioned above. The magnitude of this flux depends strongly on the cosmic-ray spectrum at acceleration, the cosmic-ray composition, and the cosmological evolution of the cosmic-ray sources. In the case that UHECR have mixed composition only the proton fraction would produce cosmogenic neutrinos. Heavy nuclei propagation produces mostly  $\bar{\nu}_e$  at lower energy from neutron decay.

The expected rate of cosmogenic neutrinos is lower than current limits obtained by IceCube [119], the Auger observatory [120], RICE [121], and ANITA-2 [122], which are shown in Fig. 29.10 together with a model for cosmogenic neutrino production [123] and the Waxman-Bahcall benchmark flux of neutrinos produced in cosmic ray sources [124]. At production, the dominant component of neutrinos



**Figure 29.10:** Differential limits on the flux of cosmogenic neutrinos set by four neutrino experiments. The curves show the Waxman-Bahcall benchmark flux (WB, [124]) and a representative midrange model for the expected flux of cosmogenic neutrinos (ESS, [123]). The expected flux is uncertain by over an order of magnitude in either direction.

comes from  $\pi^\pm$  decays and has flavor content  $\nu_e : \nu_\mu : \nu_\tau = 1 : 2 : 0$ . After oscillations, the arriving cosmogenic neutrinos are expected to be an equal mixture of all three flavors. The sensitivity of each experiment depends on neutrino flavor. IceCube, RICE, and ANITA are sensitive to all three flavors, and the sensitivity to different flavors is energy dependent. The limit of Auger is only for  $\nu_\tau$  and  $\bar{\nu}_\tau$  which should be about 1/3 of the total neutrino flux after oscillations, so this limit is plotted multiplied by a factor of three for comparison with the other limits and with the theoretical estimates.

IceCube has reported a population of neutrino events extending from 30 TeV up to 2 PeV that exceeds expected atmospheric backgrounds and is thus most likely of astrophysical origin [126,127].

#### References:

1. D. Maurin, *et al.*, [arXiv:1302.5525](#).
2. M. Boezio *et al.*, *Astropart. Phys.* **19**, 583 (2003).
3. AMS Collab., *Phys. Lett.* **B490**, 27 (2000); *Phys. Lett.* **B494**, 193 (2000).
4. T. Sanuki *et al.*, *Astrophys. J.* **545**, 1135 (2000).
5. S. Haino *et al.*, *Phys. Lett.* **B594**, 35 (2004).
6. H.S. Ahn *et al.*, *Astrophys. J.* **707**, 593 (2000).
7. J.J. Engelmann *et al.*, *Astron. & Astrophys.* **233**, 96 (1990).
8. D. Müller *et al.*, *Ap J*, **374**, 356 (1991).
9. M. Ave *et al.*, *Astrophys. J.* **678**, 262 (2008).
10. A.D. Panov *et al.*, *Bull. Russian Acad. of Science, Physics*, **71**, 494 (2007).
11. V.A. Derbina *et al.*, *Astrophys. J.* **628**, L41 (2005).
12. K. Asakimori *et al.* (JACEE Collab.), *Astrophys. J.* **502**, 278 (1998).
13. F. Aharonian *et al.* (HESS Collab.), *Phys. Rev.* **D75**, 042004 (2007).
14. A.W. Strong *et al.*, *Ann. Rev. Nucl. and Part. Sci.* **57**, 285 (2007).
15. R.A. Mewaldt *et al.*, *Space Science Reviews* **99**, 27 (2001).
16. A.A. Abdo *et al.*, *Astrophys. J.* **698**, 2121 (2009).
17. R. Abbasi *et al.*, *Astrophys. J.* **718**, L194 (2010).
18. M. Amenomori *et al.*, *Astrophys. J.* **711**, 119 (2010).
19. J. Chang *et al.* (ATIC Collab.), *Nature* **456**, 362 (2008).
20. A.A. Abdo *et al.* (Fermi/LAT Collab.), *Phys. Rev. Lett.* **102**, 181101 (2009); M. Ackermann *et al.*, *Phys. Rev.* **D82**, 092004 (2010).
21. M. Aguilar *et al.* (AMS-02 Collab.), *Phys. Rev. Lett.* **113**, 221102 (2014).
22. F. Aharonian *et al.* (HESS Collab.), *Phys. Rev. Lett.* **101**, 261104 (2008) and *Astron. & Astrophys.* **508**, 561 (2009).
23. O. Adriani *et al.*, *Phys. Rev. Lett.* **106**, 201101 (2011).
24. L. Accardo *et al.* (AMS-02 Collab.), *Phys. Rev. Lett.* **113**, 121101 (2014).
25. Y. Shizake *et al.*, *Astropart. Phys.* **28**, 154 (2007).
26. O. Adriani *et al.* (Pamela Collab.), *Nature* **458**, 607 (2009); *Phys. Rev. Lett.* **102**, 051101 (2009).
27. M. Aguilar *et al.*, *Phys. Rev. Lett.* **110**, 141102 (2013).
28. I.V. Moskalenko and A.W. Strong, *Astrophys. J.* **493**, 694 (1998).
29. A. Ibarra *et al.*, *Int. J. Mod. Phys.* **A28**, 1330040 (2013).
30. J.J. Beatty *et al.*, *Phys. Rev. Lett.* **93**, 24112 (2004).
31. J. Nishimura *et al.*, *Adv. Space Research* **19**, 767 (1997).
32. D. Gaggero *et al.*, *Phys. Rev. Lett.* **111**, 021102 (2013).
33. M.A. DuVernois *et al.*, *Astrophys. J.* **559**, 296 (2001).
34. M. Boezio *et al.*, *Astrophys. J.* **532**, 653 (2000).
35. A.S. Beach *et al.*, *Phys. Rev. Lett.* **87**, 271101 (2001).
36. A. Yamamoto *et al.*, *Adv. Space Research* **42**, 443 (2008).
37. Y. Asaoka *et al.*, *Phys. Rev. Lett.* **88**, 51101 (2002).
38. K. Abe *et al.*, *Phys. Rev. Lett.* **108**, 081102 (2013).
39. H. Fuke *et al.*, *Phys. Rev. Lett.* **95**, 081101 (2005).
40. P. Yin *et al.*, *Phys. Rev.* **D88**, 023001 (2013).
41. R. Bellotti *et al.*, *Phys. Rev.* **D53**, 35 (1996).
42. R. Bellotti *et al.*, *Phys. Rev.* **D60**, 052002 (1999).
43. M. Boezio *et al.*, *Phys. Rev.* **D62**, 032007 (2000); M. Boezio *et al.*, *Phys. Rev.* **D67**, 072003 (2003).
44. S. Coutu *et al.*, *Phys. Rev.* **D62**, 032001 (2000).
45. S. Haino *et al.*, *Phys. Lett.* **B594**, 35 (2004).
46. T. Sanuki *et al.*, *Phys. Rev.* **D75**, 043005 (2007).
47. T.K. Gaisser, *Cosmic Rays and Particle Physics*, Cambridge University Press (1990).
48. P. Lipari, *Astropart. Phys.* **1**, 195 (1993).
49. E. Mocciutto *et al.*, in *Proc. 28th Int. Cosmic Ray Conf.*, Tsukuba, 1627 (2003). [<http://adsabs.harvard.edu/abs/2003ICRC...3.1627M>].
50. M.P. De Pascale *et al.*, *J. Geophys. Res.* **98**, 3501 (1993).
51. P.K.F. Grieder, *Cosmic Rays at Earth*, Elsevier Science (2001).
52. J. Kremer *et al.*, *Phys. Rev. Lett.* **83**, 4241 (1999).
53. S. Haino *et al.* (BESS Collab.), *Phys. Lett.* **B594**, 35 (2004).
54. P. Archard *et al.* (L3+C Collab.), *Phys. Lett.* **B598**, 15 (2004).
55. C.G.S. Costa, *Astropart. Phys.* **16**, 193 (2001).
56. O.C. Allkofer, K. Carstensen, and W.D. Dau, *Phys. Lett.* **B36**, 425 (1971).
57. B.C. Rastin, *J. Phys.* **G10**, 1609 (1984).
58. C.A. Ayre *et al.*, *J. Phys.* **G1**, 584 (1975).
59. H. Jokisch *et al.*, *Phys. Rev.* **D19**, 1368 (1979).
60. P. Adamson *et al.* (MINOS Collab.), *Phys. Rev.* **D76**, 052003 (2007).
61. S. Hayakawa, *Cosmic Ray Physics*, Wiley, Interscience, New York (1969).
62. R.R. Daniel and S.A. Stephens, *Revs. Geophysics & Space Sci.* **12**, 233 (1974).
63. K.P. Beuermann and G. Wibberenz, *Can. J. Phys.* **46**, S1034 (1968).
64. I.S. Diggory *et al.*, *J. Phys.* **A7**, 741 (1974).
65. V. Khachatryan *et al.* (CMS Collab.) *Phys. Lett.* **B692**, 83 (2010).
66. N. Agafonova *et al.* (OPERA Collab.) *Eur. Phys. J.* **C67**, 25 (2010).
67. D.E. Groom, N.V. Mokhov, and S.I. Striganov, "Muon stopping-power and range tables," *Atomic Data and Nuclear Data Tables*, **78**, 183 (2001).
68. P. Lipari and T. Stanev, *Phys. Rev.* **D44**, 3543 (1991).

69. M. Crouch, in *Proc. 20th Int. Cosmic Ray Conf.*, Moscow, **6**, 165 (1987)  
[<http://adsabs.harvard.edu/abs/1987ICRC....6..165C>].
70. I.A. Belolaptikov *et al.*, *Astropart. Phys.* **7**, 263 (1997).
71. J. Babson *et al.*, *Phys. Rev.* **D42**, 3613 (1990).
72. P. Desiati *et al.*, in *Proc. 28th Int. Cosmic Ray Conf.*, Tsukuba, 1373 (2003)  
[<http://adsabs.harvard.edu/abs/2003ICRC....3.1373D>].
73. T. Pradier *et al.* (ANTARES Collab.), [arXiv:0805.2545](https://arxiv.org/abs/0805.2545) and 31st ICRC, 7-15 July 2009, Łódź, Poland (paper #0340).
74. Yu.M. Andreev, V.I. Gurentzov, and I.M. Kogai, in *Proc. 20th Int. Cosmic Ray Conf.*, Moscow, **6**, 200 (1987),  
[<http://adsabs.harvard.edu/abs/1987ICRC....6..200A>].
75. M. Aglietta *et al.* (LVD Collab.), *Astropart. Phys.* **3**, 311 (1995).
76. M. Ambrosio *et al.* (MACRO Collab.), *Phys. Rev.* **D52**, 3793 (1995).
77. Ch. Berger *et al.* (Frejus Collab.), *Phys. Rev.* **D40**, 2163 (1989).
78. C. Waltham *et al.*, in *Proc. 27th Int. Cosmic Ray Conf.*, Hamburg, 991 (2001),  
[<http://adsabs.harvard.edu/abs/2001ICRC....3..991W>].
79. Y. Ashie *et al.* (SuperKamiokande Collab.), *Phys. Rev.* **D71**, 112005 (2005).
80. T. Futagami *et al.*, *Phys. Rev. Lett.* **82**, 5194 (1999).
81. Y. Fukuda *et al.*, *Phys. Rev. Lett.* **81**, 1562 (1998).
82. F. Reines *et al.*, *Phys. Rev. Lett.* **15**, 429 (1965).
83. M.M. Boliev *et al.*, in *Proc. 3rd Int. Workshop on Neutrino Telescopes* (ed. Milla Baldo Ceolin), 235 (1991).
84. M. Ambrosio *et al.*, (MACRO) *Phys. Lett.* **B434**, 451 (1998).  
The number quoted for MACRO is the average over 90% of the lower hemisphere,  $\cos\theta < -0.1$ ; see F. Ronga *et al.*, [hep-ex/9905025](https://arxiv.org/abs/hep-ex/9905025).
85. R. Becker-Szendy *et al.*, *Phys. Rev. Lett.* **69**, 1010 (1992);  
*Proc. 25th Int. Conf. High-Energy Physics*, Singapore (eds. K.K. Phua and Y. Yamaguchi, World Scientific), 662 1991.
86. S. Hatakeyama *et al.*, *Phys. Rev. Lett.* **81**, 2016 (1998).
87. Y. Fukuda *et al.*, *Phys. Rev. Lett.* **82**, 2644 (1999).
88. K. Greisen, *Ann. Rev. Nucl. Sci.* **10**, 63 (1960).
89. S.P. Swordy *et al.*, *Astropart. Phys.* **18**, 129 (2002).
90. N.L. Grigorov *et al.*, *Sov. J. Nucl. Phys.* **11**, 588. *Proc. 12th Int. Cosmic Ray Conf.*, Hobart, **1**, 1746 and 1752 (1971).
91. K. Asakimori *et al.*, *Proc. 23rd Int. Cosmic Ray Conf.*, Calgary, **2**, 25 (1993);  
*Proc. 22nd Int. Cosmic Ray Conf.*, Dublin, **2**, 57 and 97 (1991)  
[<http://adsabs.harvard.edu/abs/1991ICRC....2...57A>]  
[<http://adsabs.harvard.edu/abs/1991ICRC....2...97A>].
92. T.V. Danilova *et al.*, *Proc. 15th Int. Cosmic Ray Conf.*, Plovdiv, **8**, 129 (1977)  
[<http://adsabs.harvard.edu/abs/1977ICRC....8..129D>].
93. Yu. A. Fomin *et al.*, *Proc. 22nd Int. Cosmic Ray Conf.*, Dublin, **2**, 85 (1991)  
[<http://adsabs.harvard.edu/abs/1991ICRC....2...85F>].
94. M. Amenomori *et al.*, *Astrophys. J.* **461**, 408 (1996).
95. M. Nagano *et al.*, *J. Phys.* **G10**, 1295 (1984).
96. F. Arquerros *et al.*, *Astron. & Astrophys.* **359**, 682 (2000).
97. M.A.K. Glasmacher *et al.*, *Astropart. Phys.* **10**, 291 (1999).
98. T. Antoni *et al.* (Kascade Collab.), *Astropart. Phys.* **24**, 1 (2005).
99. M. Amenomori *et al.*, *Astrophys. J.* **268**, 1165 (2008).
100. W.D. Apel *et al.*, *Phys. Rev. Lett.* **107**, 171104 (2011).
101. M.G. Aartsen *et al.*, (IceCube Collab.) [arXiv:1307.3795v1](https://arxiv.org/abs/1307.3795v1)(2013).
102. R. Abbasi *et al.*, (HiRes Collab.), *Phys. Rev. Lett.* **100**, 101101 (2008).
103. J. Abraham *et al.*, (Auger Collab.), *Phys. Rev. Lett.* **101**, 061101 (2008).
104. D. Ivanov *et al.*, (Telescope Array Collab.), *Proceedings of Science* (ICRC2015), 349 (2015).
105. I. Valiño *et al.*, (Auger Collab.), *Proceedings of Science* (ICRC2015), 271 (2015).
106. D.J. Bird *et al.*, (Fly's Eye Collab.), *Astrophys. J.* **424**, 491 (1994).
107. M. Takeda *et al.*, (The AGASA Collab.), *Astropart. Phys.* **19**, 447 (2003).
108. P. Tinyakov *et al.*, (Telescope Array Collab.), *Proceedings of Science* (ICRC2015), 326 (2015).
109. V.S. Ptuskin *et al.*, *Astron. & Astrophys.* **268**, 726 (1993).
110. V.S. Berezinsky and S.I. Grigor'eva, *Astron. & Astrophys.* **199**, 1 (1988).
111. V. Berezinsky, A. Gazizov, and S. Grigorieva, *Phys. Rev.* **D74**, 043005 (2006).
112. R.U. Abbasi *et al.*, (The HiRes Collab.), *Astrophys. J.* **622**, 910 (2005).
113. M. Unger *et al.*, (Auger Collab.), *Proc. 30th Int. Cosmic Ray Conf.*, Merida, Mexico, 2007 ([arXiv:0706.1495](https://arxiv.org/abs/0706.1495)).
114. J. Abraham *et al.*, (Auger Collab.), *Proc. 31st Int. Cosmic Ray Conf.*, Lodz, Poland, 2009; ([arXiv:0906.2319](https://arxiv.org/abs/0906.2319)).
115. K. Greisen, *Phys. Rev. Lett.* **16**, 748 (1966).
116. G.T. Zatsepin and V.A. Kuz'min, *Sov. Phys. JETP Lett.* **4**, 78 (1966).
117. D. Allard *et al.*, *Astron. & Astrophys.* **443**, L29 (2005).
118. V.S. Berezinsky and G.T. Zatsepin, *Phys. Lett.* **B28**, 423 (1969).
119. M.G. Aartsen *et al.*, (IceCube Collab.), *Phys. Rev.* **D88**, 112008 (2013).
120. A. Aab *et al.*, (Auger Collab.), *Phys. Rev.* **D91**, 092008 (2015).
121. I. Kravchenko *et al.*, (RICE Collab.), *Phys. Rev.* **D73**, 082002 (2006);  
I. Kravchenko *et al.*, [arXiv:1106.1164](https://arxiv.org/abs/1106.1164) (2011).
122. P. Gorham *et al.*, (ANITA Collab.), *Phys. Rev.* **D82**, 022004 (2010);  
P. Gorham *et al.*, (ANITA Collab.), ([arXiv:1011.5004](https://arxiv.org/abs/1011.5004)).
123. R. Engel, D. Seckel, and T. Stanev, *Phys. Rev.* **D64**, 09310 (2001).
124. E. Waxman and J. Bahcall, *Phys. Rev.* **D59**, 023002 (1999).
125. R.U. Abbasi *et al.*, (HiRes Collab.), *Astropart. Phys.* **32**, 53 (2009).
126. M. Aartsen *et al.*, (IceCube Collab.) *Science* **342**, 1242856 (2013).
127. M. Aartsen *et al.*, (IceCube Collab.) *Phys. Rev. Lett.* **113**, 101101 (2013).

### 30. ACCELERATOR PHYSICS OF COLLIDERS

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#### 30.1. Luminosity

This article provides background for the High-Energy Collider Parameter Tables that follow. The number of events,  $N_{exp}$ , is the product of the cross section of interest,  $\sigma_{exp}$ , and the time integral over the instantaneous *luminosity*,  $\mathcal{L}$ :

$$N_{exp} = \sigma_{exp} \times \int \mathcal{L}(t) dt. \quad (30.1)$$

Today's colliders all employ bunched beams. If two bunches containing  $n_1$  and  $n_2$  particles collide head-on with frequency  $f_{coll}$ , a basic expression for the luminosity is

$$\mathcal{L} = f_{coll} \frac{n_1 n_2}{4\pi\sigma_x\sigma_y} \quad (30.2)$$

where  $\sigma_x$  and  $\sigma_y$  characterize the rms transverse beam sizes in the horizontal (bend) and vertical directions. In this form it is assumed that the bunches are identical in transverse profile, that the profiles are Gaussian and independent of position along the bunch, and the particle distributions are not altered during bunch crossing. Nonzero beam crossing angles and long bunches will reduce the luminosity from this value.

Whatever the distribution at the source, by the time the beam reaches high energy, the normal form is a useful approximation as suggested by the  $\sigma$ -notation. In the case of an electron storage ring, synchrotron radiation leads to a Gaussian distribution in equilibrium, but even in the absence of radiation the central limit theorem of probability and the diminished importance of space charge effects produce a similar result.

The luminosity may be obtained directly by measurement of the beam properties in Eq. (30.2). For continuous measurements, an expression similar to Eq. (30.1) with  $N_{ref}$  from a known reference cross section,  $\sigma_{ref}$ , may be used to determine  $\sigma_{exp}$  according to  $\sigma_{exp} = (N_{exp}/N_{ref})\sigma_{ref}$ .

In the Tables, luminosity is stated in units of  $\text{cm}^{-2}\text{s}^{-1}$ . Integrated luminosity, on the other hand is usually quoted as the inverse of the standard measures of cross section such as femtobarns and, recently, attobarns. Subsequent sections in this report briefly expand on the dynamics behind collider design, comment on the realization of collider performance in a selection of today's facilities, and end with some remarks on future possibilities.

#### 30.2. Beam Dynamics

The first concern of beam dynamics is stability. While a reference particle proceeds along the design, or reference, trajectory other particles in the bunch are to remain close by. Assume that the reference particle carries a right-handed Cartesian coordinate system, with the  $z$ -coordinate pointed in the direction of motion along the reference trajectory. The independent variable is the distance  $s$  of the reference particle along this trajectory rather than time, and for simplicity this path is taken to be planar. The transverse coordinates are  $x$  and  $y$ , where  $\{x, z\}$  defines the plane of the reference trajectory. Several time scales are involved, and the approximations used in writing the equations of motion reflect that circumstance. All of today's high energy colliders are alternating-gradient synchrotrons or, respectively, storage rings [1,2], and the shortest time scale is that associated with transverse motion, that is described in terms of betatron oscillations, so called because of their analysis for the betatron accelerator species years ago. The linearized equations of motion of a particle displaced from the reference particle are

$$\begin{aligned} x'' + K_x x &= 0, & K_x &\equiv \frac{q}{p} \frac{\partial B}{\partial x} + \frac{1}{\rho^2} \\ y'' + K_y y &= 0, & K_y &\equiv -\frac{q}{p} \frac{\partial B}{\partial y} \\ z' &= -x/\rho \end{aligned} \quad (30.3)$$

where the magnetic field  $B(s)$  along the design trajectory is only in the  $y$  direction, contains only dipole and quadrupole terms, and is treated as static here. The radius of curvature due to the field on the reference orbit is  $\rho$ ;  $z$  represents the longitudinal distance from the reference particle;  $p$  and  $q$  are the particle's momentum and charge, respectively. The prime denotes  $d/ds$ . The pair  $(x, x')$  describes approximately-canonical variables. For more general cases (e.g. acceleration) one should use  $(x, p_x)$  instead, where  $p_x$  denotes the transverse momentum in the  $x$ -direction.

The equations for  $x$  and  $y$  are those of harmonic oscillators but with a restoring force periodic in  $s$ ; that is, they are instances of Hill's equation. The solution may be written in the form

$$\begin{aligned} x(s) &= A_x \sqrt{\beta_x} \cos \psi_x \\ x'(s) &= -\frac{A_x}{\sqrt{\beta_x}} [\alpha_x \cos \psi_x + \sin \psi_x] \end{aligned} \quad (30.4)$$

where  $A_x$  is a constant of integration,  $\alpha_x \equiv -(1/2)d\beta_x(s)/ds$ , and the envelope of the motion is modulated by the *amplitude function*,  $\beta_x$ . A solution of the same form describes the motion in  $y$ . The subscripts will be suppressed in the following discussion.

The amplitude function satisfies

$$2\beta\beta'' - \beta'^2 + 4\beta^2 K = 4, \quad (30.5)$$

and in a region free of magnetic field it should be noted that the solution of Eq. (30.5) is a parabola. Expressing  $A$  in terms of  $x, x'$  yields

$$\begin{aligned} A^2 &= \gamma x^2 + 2\alpha x x' + \beta x'^2 \\ &= \frac{1}{\beta} [x^2 + (\alpha x + \beta x')^2] \end{aligned} \quad (30.6)$$

with  $\gamma \equiv (1 + \alpha^2)/\beta$ . In a single pass system such as a linac, the *Courant-Snyder parameters*  $\alpha, \beta, \gamma$  may be selected to match the  $x, x'$  distribution of the input beam; in a recursive system, the parameters are usually defined by the structure rather than by the beam.

The relationships between the parameters and the structure may be seen by treatment of a simple *lattice* consisting of equally-spaced thin-lens quadrupoles whose magnetic-field gradients are equal in magnitude but alternating in sign. For this discussion, the weak focusing effects of the bending magnets may be neglected. The propagation of  $X \equiv \{x, x'\}$  through a repetition period may be written  $X_2 = M X_1$ , with the matrix  $M = FODO$  composed of the matrices

$$F = \begin{pmatrix} 1 & 0 \\ -1/f & 1 \end{pmatrix}, \quad D = \begin{pmatrix} 1 & 0 \\ 1/f & 1 \end{pmatrix}, \quad O = \begin{pmatrix} 1 & L \\ 0 & 1 \end{pmatrix},$$

where  $f$  is the magnitude of the focal length and  $L$  the lens spacing. Then

$$M = \begin{pmatrix} 1 + \frac{L}{f} & 2L + \frac{L^2}{f} \\ -\frac{L}{f^2} & 1 - \frac{L}{f} - \frac{L^2}{f^2} \end{pmatrix}. \quad (30.7)$$

The matrix for  $y$  is identical in form differing only by a change in sign of the terms linear in  $1/f$ . An eigenvector-eigenvalue analysis of the matrix  $M$  shows that the motion is stable provided  $f > L/2$ . While that criterion is easily met, in practice instability may be caused by many other factors, including the beam-beam interaction itself.

Standard focus-drift-defocus-drift, or *FODO*, cells such as characterized in simple form by Eq. (30.7) occupy most of the layout of a large collider ring and may be used to set the scale of the amplitude function and related phase advance. Conversion of Eq. (30.4) to a matrix form equivalent to Eq. (30.7) (but more generally valid, i.e. for any stable periodic linear motion) gives

$$M = \begin{pmatrix} C + \alpha S & \beta S \\ -\gamma S & C - \alpha S \end{pmatrix} \quad (30.8)$$

where  $C \equiv \cos \Delta\psi$ ,  $S \equiv \sin \Delta\psi$ , and the relation between structure and amplitude function is specified by setting the values of the

latter to be the same at both ends of the cell. By comparison of Eq. (30.7) and Eq. (30.8) one finds  $C = 1 - L^2/(2f^2)$ , so that the choice  $f = L/\sqrt{2}$  would give a phase advance  $\Delta\psi$  of 90 degrees for the standard cell. The amplitude function would have a maximum at the focusing quadrupole of magnitude  $\beta = 2.7L$ , illustrating the relationship of alternating gradient focusing amplitudes to relatively local aspects of the design. Other functionalities such as injection, extraction, and HEP experiments are included by lattice sections matched to the standard cell parameters ( $\beta$ ,  $\alpha$ ) at the insertion points. The phase advances according to  $d\psi/ds = 1/\beta$ ; that is,  $\beta$  also plays the role of a local  $\lambda/2\pi$ , and the *tune*,  $\nu$ , is the number of such oscillations per turn about the closed path. In the neighborhood of an interaction point (IP), the beam optics of the ring is configured so as to produce a narrow focus; the value of the amplitude function at this point is designated  $\beta^*$ .

The motion as it develops with  $s$  describes an ellipse in  $\{x, x' \equiv dx/ds\}$  phase space, the area of which is  $\pi A^2$ , where  $A$  is the constant in Eq. (30.4). If the interior of that ellipse is populated by an ensemble of non-interacting particles, that area, given the name *emittance* and denoted by  $\varepsilon$ , would change only with energy. More precisely, for a beam with a Gaussian distribution in  $x, x'$ , the area containing one standard deviation  $\sigma_x$ , divided by  $\pi$ , is used as the definition of emittance in the Tables:

$$\varepsilon_x \equiv \frac{\sigma_x^2}{\beta_x}, \quad (30.9)$$

with a corresponding expression in the other transverse direction,  $y$ . This definition includes 39% of the beam. For most of the entries in the Tables the standard deviation is used as the beam radius.

To complete the coordinates used to describe the motion, we take as the variable conjugate to  $z$  the fractional momentum deviation  $\delta p/p$  from that of the reference particle. Radiofrequency electric fields in the  $s$  direction provide a means for longitudinal oscillations, and the frequency determines the bunch length. The frequency of this system appears in the Tables as does the rms value of  $\delta p/p$  characterized as “energy spread” of the beam.

For HEP bunch length is a significant quantity for a variety of reasons, but in the present context if the bunch length becomes larger than  $\beta^*$  the luminosity is adversely affected. This is because  $\beta$  grows parabolically as one proceeds away from the interaction point and so the beam size increases thus lowering the contribution to the luminosity from such locations. This is often called the “hourglass” effect.

The other major external electromagnetic field interaction in the single particle context is the production of synchrotron radiation due to centripetal acceleration, given by the Larmor formula multiplied by a relativistic magnification factor of  $\gamma^4$  [3]. In the case of electron rings this process determines the equilibrium emittance through a balance between radiation damping and excitation of oscillations, and further serves as a barrier to future higher energy versions in this variety of collider. A related phenomenon is beamstrahlung, i.e. the synchrotron radiation emitted during the collision in the field of the opposing beam, which is relevant for both linear colliders (where it degrades the luminosity spectrum) and future highest-energy circular colliders (where it limits the beam lifetime). For both types of colliders the beamstrahlung is mitigated by making the colliding beams as flat as possible ( $\sigma_x^* \gg \sigma_y^*$ ).

A more comprehensive discussion of betatron oscillations, longitudinal motion, and synchrotron radiation is available in the 2008 version of the PDG review [4].

### 30.3. Road to High Luminosity

Eq. (30.2) can be recast in terms of emittances and amplitude functions as

$$\mathcal{L} = f \frac{n_1 n_2}{4\pi \sqrt{\varepsilon_x \beta_x^* \varepsilon_y \beta_y^*}}. \quad (30.10)$$

So to achieve high luminosity, all one has to do is make high population bunches of low emittance collide at high frequency at locations where

the beam optics provides as low values of the amplitude functions as possible.

Such expressions as Eq. (30.10) of the luminosity are special cases of the more general forms available elsewhere [5], wherein the reduction due to crossing angle and other effects can be found. But while there are no fundamental limits to the process, there are certainly challenges. Here we have space to mention only a few of these. The beam-beam tune shift appears in the Tables. A bunch in beam 1 presents a (nonlinear) lens to a particle in beam 2 resulting in changes to the particle's transverse tune with a range characterized by the parameter [5]

$$\xi_{y,2} = \frac{\mu_0}{8\pi^2} \frac{q_1 q_2 n_1 \beta_{y,2}^*}{m_{A,2} \gamma_2 \sigma_{y,1} (\sigma_{x,1} + \sigma_{y,1})} \quad (30.11)$$

where  $q_1$  ( $q_2$ ) denotes the particle charge of beam 1 (2) in units of the elementary charge,  $m_{A,2}$  the mass of beam-2 particles, and  $\mu_0$  the vacuum permeability. The transverse oscillations are susceptible to resonant perturbations from a variety of sources such as imperfections in the magnetic guide field, so that certain values of the tune must be avoided. Accordingly, the tune spread arising from  $\xi$  is limited, but limited to a value difficult to predict. But a glance at the Tables shows that electrons are more forgiving than protons thanks to the damping effects of synchrotron radiation; the  $\xi$ -values for the former are about an order of magnitude larger than those for protons.

A subject of present intense interest is the *electron-cloud effect* [6,7]; actually a variety of related processes come under this heading. They typically involve a buildup of electron density in the vacuum chamber due to emission from the chamber walls stimulated by electrons or photons originating from the beam itself. For instance, there is a process closely resembling the multipacting effects familiar from radiofrequency system commissioning. Low energy electrons are ejected from the walls by photons from positron or proton beam-produced synchrotron radiation. These electrons are accelerated toward a beam bunch, but by the time they reach the center of the vacuum chamber the bunch has gone and so the now-energetic electrons strike the opposite wall to produce more secondaries. These secondaries are now accelerated by a subsequent bunch, and so on. Among the disturbances that this electron accumulation can produce is an enhancement of the tune spread within the bunch; the near-cancellation of bunch-induced electric and magnetic fields is no longer in effect.

If the luminosity of Eq. (30.10) is rewritten in terms of the beam-beam parameter, Eq. (30.11), the emittance itself disappears. However, the emittance must be sufficiently small to realize a desired magnitude of beam-beam parameter, but once  $\xi_y$  reaches this limit, further lowering the emittance does not lead to higher luminosity.

For electron synchrotrons and storage rings, radiation damping provides an automatic route to achieve a small emittance. In fact, synchrotron radiation is of key importance in the design and optimization of  $e^+e^-$  colliders. While vacuum stability and electron clouds can be of concern in the positron rings, synchrotron radiation along with the restoration of longitudinal momentum by the RF system has the positive effect of generating very small transverse beam sizes and small momentum spread. Further reduction of beam size at the interaction points using standard beam optics techniques and successfully contending with high beam currents has led to record luminosities in these rings, exceeding those of hadron colliders. To maximize integrated luminosity the beam can be “topped off” by injecting new particles without removing existing ones – a feature difficult to imitate in hadron colliders.

For hadrons, particularly antiprotons, two inventions have played a prominent role. Stochastic cooling [8] was employed first to prepare beams for the  $S\bar{p}pS$  and subsequently in the Tevatron and in RHIC [9,10]. Electron cooling [11] was also used in the Tevatron complex to great advantage. Further innovations are underway driven by the needs of potential future projects; these are noted in the final section.

### 30.4. Recent High Energy Colliders

Collider accelerator physics of course goes far beyond the elements of the preceding sections. In this and the following section elaboration is made on various issues associated with some of the recently operating colliders, particularly factors which impact integrated luminosity. The various colliders utilizing hadrons each have unique characteristics and are, therefore, discussed separately. As space is limited, general references are provided where much further information can be obtained. A more complete list of recent colliders and their parameters can be found in the High-Energy Collider Parameters tables.

**30.4.1. *Tevatron* :** [12] The first synchrotron in history using superconducting magnets, the Tevatron, was the highest energy collider for 25 years. Operation was terminated in September 2011, after delivering more than  $10 \text{ fb}^{-1}$  to the p-p collider experiments CDF and D0. The route to high integrated luminosity in the Tevatron was governed by the antiproton production rate, the turn-around time to produce another store, and the resulting optimization of store time. The proton and antiproton beams in the Tevatron circulated in a single vacuum pipe and thus were placed on separated orbits which wrapped around each other in a helical pattern outside of the interaction regions. Hence, long-range encounters played an important role here as well, with the 70 long-range encounters distributed about the synchrotron, and mitigation was limited by the available aperture. The Tevatron ultimately achieved luminosities a factor of 400 over its original design specification.

**30.4.2. *HERA* :** [13] HERA, operated between 1992 and 2007, delivered nearly  $1 \text{ fb}^{-1}$  of integrated luminosity to the electron-proton collider experiments H1 and ZEUS. HERA was the first high-energy lepton-hadron collider, and also the first facility to employ both applications of superconductivity: magnets and accelerating structures. The proton beams of HERA had a maximum energy of 920 GeV. The lepton beams (positrons or electrons) were provided by the existing DESY complex, and were accelerated to 27.5 GeV using conventional magnets. At collision a 4-times higher frequency RF system, compared with the injection RF, was used to generate shorter bunches, thus helping alleviate the hourglass effect at the collision points. The lepton beam naturally would become transversely polarized (within about 40 minutes) and “spin rotators” were implemented on either side of an IP to produce longitudinal polarization at the experiment.

**30.4.3. *LEP* :** [14] Installed in a tunnel of 27 km circumference, LEP was the largest circular  $e^+e^-$  collider built so far. It was operated from 1989 to 2000 with beam energies ranging from 45.6 to 104.5 GeV and a maximum luminosity of  $10^{32} \text{ cm}^{-2}\text{s}^{-1}$ , at 98 GeV, surpassing all relevant design parameters.

**30.4.4. *SLC* :** [15] Based on an existing 3-km long S-band linac, the SLC was the first and only linear collider. It was operated from 1987 to 1998 with a constant beam energy of 45.6 GeV, up to about 80% electron-beam polarization, quasi-flat beams, and, in its last year, a typical peak luminosity of  $2 \times 10^{30} \text{ cm}^{-2}\text{s}^{-1}$ , a third of the design value.

### 30.5. Present Collider Facilities

**30.5.1. *LHC* :** [16] The superconducting Large Hadron Collider is the world’s highest energy collider. In 2012 operation for HEP has been at 4 TeV per proton [17]. The beam energy increased to 6.5 TeV in 2015. The current status is best checked at the Web site [18]. To meet its luminosity goals the LHC will have to contend with a high beam current of 0.5 A, leading to stored energies of several hundred MJ per beam. Component protection, beam collimation, and controlled energy deposition are given very high priorities. Additionally, at energies of 5-7 TeV per particle, synchrotron radiation

will move from being a curiosity to a challenge in a hadron accelerator for the first time. At design beam current the cryogenic system must remove roughly 7 kW due to synchrotron radiation, intercepted at a temperature of 4.5-20 K. As the photons are emitted their interactions with the vacuum chamber wall can generate free electrons, with consequent “electron cloud” development. Much care was taken to design a special beam screen for the chamber to mitigate this issue.

The two proton beams are contained in separate pipes throughout most of the circumference, and are brought together into a single pipe at the interaction points. The large number of bunches, and subsequent short bunch spacing, would lead to approximately 30 head-on collisions through 120 m of common beam pipe at each IP. Thus, a small crossing angle is employed, which reduces the luminosity by about 15%. Still, the bunches moving in one direction will have long-range encounters with the counter-rotating bunches and the resulting perturbations of the particle motion constitute a continued course of study. The luminosity scale is absolutely calibrated by the “van der Meer method” as was invented for the ISR [19], and followed by multiple, redundant luminosity monitors (see for example [20] and references therein). The Tables also show the performance anticipated for Pb-Pb collisions. The ALICE [21] experiment is designed to concentrate on these high energy-density phenomena, which are studied as well by ATLAS and CMS. The LHC can also provide Pb-p collisions as it did in early 2013.

In the coming years, an ambitious upgrade program, HL-LHC [22], has as its target an order-of-magnitude increase in luminosity through the utilization of Nb<sub>3</sub>Sn superconducting magnets, superconducting compact “crab” cavities and luminosity leveling as key ingredients.

**30.5.2.  $e^+e^-$  *Rings* :** Asymmetric energies of the two beams have allowed for the enhancement of B-physics research and for interesting interaction region designs. As the bunch spacing can be quite short, the lepton beams sometimes pass through each other at an angle and hence have reduced luminosity. Recently, however, the use of high frequency “crab crossing” schemes has produced full restoration of the luminous region. KEK-B attained over  $1 \text{ fb}^{-1}$  of integrated luminosity in a single day, and its upgrade, SuperKEKB, is aiming for luminosities of  $8 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$  [23]. A different collision approach, called “crab waist”, which relies on special sextupoles together with a large crossing angle, has been successfully implemented at DAΦNE [24]. Other  $e^+e^-$  ring colliders in operation are BEPC-II, VEPP-2000 and VEPP-4M [23].

**30.5.3. *RHIC* :** [25] The Relativistic Heavy Ion Collider employs superconducting magnets, and collides combinations of fully-stripped ions such as H-H (p-p), U-U, Au-Au, Cu-Au, Cu-Cu, and d-Au. The high charge per particle (+79 for gold, for instance) makes intra-beam scattering of particles within the bunch a special concern, even for seemingly moderate bunch intensities. In 2012, 3-D stochastic cooling was successfully implemented in RHIC, reducing the transverse emittances of heavy ion beams by a factor of 5 [10]. Another special feature of accelerating heavy ions in RHIC is that the beams experience a “transition energy” during acceleration – a point where the derivative with respect to momentum of the revolution period is zero. This is more typical of low-energy accelerators, where the necessary phase jump required of the RF system is implemented rapidly and little time is spent near this condition. In the case of RHIC with heavy ions, the superconducting magnets do not ramp very quickly and the period of time spent crossing transition is long and must be dealt with carefully. For p-p operation the beams are always above their transition energy and so this condition is completely avoided.

RHIC is also distinctive in its ability to accelerate and collide polarized proton beams. As proton beam polarization must be maintained from its low-energy source, successful acceleration through the myriad of depolarizing resonance conditions in high energy circular accelerators has taken years to accomplish. An energy of 255 GeV per proton with > 50% final polarization per beam has been realized. As part of a scheme to compensate the head-on beam-beam effect, electron lenses operated routinely during the polarized proton operation in 2015.



**Table 30.1:** Tentative parameters of selected future high-energy colliders. Parameters of HL-LHC, ILC and CLIC can be found in the High-Energy Collider Parameters tables.

	LHeC	FCC-ee			CEPC	FCC-hh	SPPC	$\mu$ collider
Species	$ep$	$e^+e^-$			$e^+e^-$	$pp$	$pp$	$\mu^+\mu^-$
Beam Energy (TeV)	0.06( $e$ ), 7 ( $p$ )	0.046	0.120	0.175	0.120	50	35	0.063
Circumference (km)	9( $e$ ), 27 ( $p$ )	100			54	100	54	0.3
Interaction regions	1	2			2	2 (4)	2	1
Estimated integrated luminosity per exp. ( $\text{ab}^{-1}/\text{year}$ )	0.1	10	1.0	0.2	0.25	0.2–1.0	0.5	0.001
Peak luminosity ( $10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ )	1	100	9	2	2	5–29	12	0.008
Time between collisions ( $\mu\text{s}$ )	0.025	0.005	0.6	6.0	3.6	0.025	0.025	1
Energy spread (rms, $10^{-3}$ )	0.03 ( $e$ ), 0.1( $p$ )	1.3	1.7	2.5	1.6	0.1	0.2	0.04
Bunch length (rms, mm)	0.06 ( $e$ ), 75.5( $p$ )	3.3	2.6	2.8	2.7	80	75.5	63
IP beam size ( $\mu\text{m}$ )	4.1 (round)	8(H), 0.03(V)	22(H), 0.03(V)	31(H), 0.05(V)	70(H), 0.15(V)	6.8 (inj.)	9.0 (inj.)	75
Injection energy (GeV)	1( $e$ ), 450( $p$ )	on energy (topping off)			on energy (topping off)	~3000	2100	on energy (topping off)
Transverse emittance (rms, nm)	0.43( $e$ ), 0.34( $p$ )	0.13(H), 0.001(V)	1.0(H), 0.001(V)	2.0(H), 0.002(V)	6.1(H), 0.02(V)	0.04 (inj.)	0.11	335
$\beta^*$ , amplitude function at interaction point (cm)	4.7( $e$ ), 5.0( $p$ )	50(H), 0.1(V)	50(H), 0.1(V)	50(H), 0.1(V)	80(H), 0.12(V)	110–30	75	1.7
Beam-beam tune shift per crossing ( $10^{-3}$ )	–( $e$ ), 0.4( $p$ )	140	100	90	118	5–15	6	20
RF frequency (MHz)	800( $e$ ), 400( $p$ )	400			650	400	400/200	805
Particles per bunch ( $10^{10}$ )	0.25( $e$ ), 22( $p$ )	50	10	20	38	10	20	400
Bunches per beam	–( $e$ ), 2808	60000	625	60	50	10600	5798	1
Average beam current (mA)	16( $e$ ), 883( $p$ )	1450	30	6.6	16.6	500	1000	640
Length of standard cell (m)	52.4( $e$ arc), 107( $p$ )	50	50	50	47	213	148	N/A
Phase advance per cell (deg)	310( $e$ H), 90( $e$ V), 90( $p$ )	90(H), 60(V)			60	90	90	N/A
Peak magnetic field (T)	0.264( $e$ ), 8.33( $p$ )	0.01	0.03	0.05	0.07	16	20	10
Polarization (%)	90( $e$ ), 0( $p$ )	$\geq 10$	0	0	0	0	0	0
SR power loss/beam (MW)	30( $e$ ), 0.01( $p$ )	50			52	2.4	2.1	$3 \times 10^{-5}$
Novel technology	high-energy ERL	—			—	16 T Nb <sub>3</sub> Sn magnets	20 T HTS magnets	ioniz. cool. high-p. target

### 30.6. Future High Energy Colliders and Prospects

Recent accomplishments of particle physics have been obtained through high-energy and high-intensity experiments using hadron-hadron, lepton-lepton, and lepton-proton colliders. Following the discovery of the Higgs particle at the LHC and in view of ongoing searches for “new physics” and rare phenomena, various options are under discussions and development to pursue future particle-physics research at higher energy and with appropriate luminosity. This is the basis for various new projects, ideas, and R&D activities, which can only briefly be summarized here. Specifically, the following projects are noted: two approaches to an electron-positron linear collider, a larger 100-km circular tunnel supporting  $e^+e^-$  collisions up to 350 or 500 GeV in the centre of mass along with a 100-TeV proton-proton collider, a muon ring collider, and potential use of plasma acceleration and other advanced schemes. Complementary studies are ongoing of a high-energy lepton-hadron collider bringing into collision a 60-GeV electron beam from an energy-recovery linac with the 7-TeV protons circulating in the LHC (LHeC) [26,27], and of  $\gamma\gamma$  collider Higgs factories based on recirculating electron linacs (e.g. SAPHIRE at CERN [28], HFiTT at FNAL [29]). Tentative parameters of some of the colliders discussed, or mentioned, in this section are summarized in Table 30.1.

**30.6.1. Electron-Positron Linear Colliders :** For three decades efforts have been devoted to develop high-gradient technology  $e^+e^-$  colliders in order to overcome the synchrotron radiation limitations of circular  $e^+e^-$  machines in the TeV energy range.

The primary challenge confronting a high energy, high luminosity single pass collider design is the power requirement, so that measures

must be taken to keep the demand within bounds as illustrated in a transformed Eq. (30.2) [30]:

$$\mathcal{L} \approx \frac{137}{8\pi r_e} \frac{P_{\text{wall}}}{E_{\text{cm}}} \frac{\eta}{\sigma_y^*} N_\gamma H_D. \quad (30.12)$$

Here,  $P_{\text{wall}}$  is the total wall-plug power of the collider,  $\eta \equiv P_b/P_{\text{wall}}$  the efficiency of converting wall-plug power into beam power  $P_b = f_{\text{coll}} n E_{\text{cm}}$ ,  $E_{\text{cm}}$  the cms energy,  $n$  ( $= n_1 = n_2$ ) the bunch population, and  $\sigma_y^*$  the vertical rms beam size at the collision point.

In formulating Eq. (30.12) the number of beamstrahlung photons emitted per  $e^\pm$ , was approximated as  $N_\gamma \approx 2\alpha r_e n / \sigma_x^*$ , where  $\alpha$  denotes the fine-structure constant. The management of  $P_{\text{wall}}$  leads to an upward push on the bunch population  $n$  with an attendant rise in the energy radiated due to the electromagnetic field of one bunch acting on the particles of the other. Keeping a significant fraction of the luminosity close to the nominal energy represents a design goal, which is met if  $N_\gamma$  does not exceed a value of about 1. A consequence is the use of flat beams, where  $N_\gamma$  is managed by the beam width, and luminosity adjusted by the beam height, thus the explicit appearance of the vertical beam size  $\sigma_y^*$ . The final factor in Eq. (30.12),  $H_D$ , represents the enhancement of luminosity due to the pinch effect during bunch crossing (the effect of which has been neglected in the expression for  $N_\gamma$ ).

The approach designated by the International Linear Collider (ILC) is presented in the Tables, and the contrast with the collision-point parameters of the circular colliders is striking, though reminiscent in direction of those of the SLAC Linear Collider. The ILC *Technical Design Report* [31] has a baseline cms energy of 500 GeV with upgrade



provision for 1 TeV, and luminosity comparable to the LHC. The ILC is based on superconducting accelerating structures of the 1.3 GHz TESLA variety.

At CERN, a design effort is underway on the Compact Linear Collider (CLIC), each linac of which is itself a two-beam accelerator, in that a high energy, low current beam is fed by a low energy, high current driver [32]. The CLIC design employs normal conducting 12 GHz accelerating structures at a gradient of 100 MeV/m, some three times the current capability of the superconducting ILC cavities. The design cms energy is 3 TeV.

**30.6.2. Future Circular Colliders :** The discovery, in 2012, of the Higgs boson at the LHC has stimulated interest in constructing a large circular tunnel which could host a variety of energy-frontier machines, including high-energy electron-positron, proton-proton, and lepton-hadron colliders. Such projects are under study by a global collaboration hosted at CERN (FCC) [33] and another one centered in China (CEPC/SPPC) [34], following earlier proposals for a Very Large Hadron Collider (VLHC) [35] and a Very Large Lepton Collider (VLLC) in the US, which would have been housed in the same 230-km long tunnel.

The maximum beam energy of a hadron collider is directly proportional to the magnetic field and to the ring circumference. The LHC magnets, based on Nb-Ti superconductor, achieve a maximum operational field of 8.33 T. The HL-LHC project develops the technology of higher field Nb<sub>3</sub>Sn magnets as well as cables made from high-temperature superconductor (HTS). Nb<sub>3</sub>Sn dipoles could ultimately reach an operational field around 15 T, and HTS inserts, requiring new engineering materials and substantial dedicated R&D, could boost this further. A cost-effective hybrid magnet design incorporating Nb-Ti, two types of Nb<sub>3</sub>Sn, and an inner layer of HTS could provide a field of 20 T [36]. If installed in the LHC tunnel, such dipoles would increase the beam energy by a factor 2.5 compared with the LHC. The vacuum system for such a machine has not yet been designed. Warm photon absorbers installed in the magnet interconnections are one of the proposed approaches, requiring experimental tests for design validation.

Further substantial increases in collision energy are possible only with a larger tunnel. The FCC hadron collider (FCC-hh) [37], formerly called VHE-LHC [38], is based on a new tunnel of about 100 km circumference, which would allow exploring energies up to 100 TeV in the centre of mass with proton-proton collisions, using 16 T magnets. This new tunnel could also accommodate a high-luminosity circular  $e^+e^-$  Higgs factory (FCC-ee) as well as a lepton-hadron collider (FCC-he).

In order to serve as a Higgs factory a new circular  $e^+e^-$  collider needs to achieve a cms energy of at least 240 GeV. FCC-ee (formerly TLEP [39]), installed in the ~100 km tunnel of the FCC-hh, could reach even higher energies, e.g. 350 GeV cms for  $t\bar{t}$  production, or up to 500 GeV for  $ZHH$  and  $Ht\bar{t}$  physics. At these energies, the luminosity, limited by the synchrotron radiation power, would still be close to  $10^{34} \text{ cm}^{-2}\text{s}^{-1}$  at each of four collision points. At lower energies (Z pole and WW threshold) FCC-ee could deliver up to three orders of magnitude higher luminosities, and also profit from radiative self polarization for precise energy calibration. The short beam lifetime at the high target luminosity, due to radiative Bhabha scattering, requires FCC-ee to be constructed as a double ring, where the collider ring operating at constant energy is complemented by a second injector ring installed in the same tunnel to “top off” the collider current. Beamstrahlung, i.e. synchrotron radiation emitted during the collision in the field of the opposing beam, introduces an additional beam lifetime limitation depending on momentum acceptance (so that achieving sufficient off-momentum dynamic aperture becomes one of the design challenges), as well as some bunch lengthening.

**30.6.3. Muon Collider :** The muon to electron mass ratio of 210 implies less concern about synchrotron radiation by a factor of about  $2 \times 10^9$  and its 2.2  $\mu\text{s}$  lifetime means that it will last for some 150B turns in a ring about half of which is occupied by bend magnets with average field  $B$  (Tesla). Design effort became serious in the mid 1990s and a collider outline emerged quickly.

Removal of the synchrotron radiation barrier reduces the scale of a muon collider facility to a level compatible with on-site placement at existing accelerator laboratories. The Higgs production cross section in the s-channel is enhanced by a factor of  $(m_\mu/m_e)^2$  compared to that in  $e^+e^-$  collisions. And a neutrino factory could potentially be realized in the course of construction [40].

The challenges to luminosity achievement are clear and amenable to immediate study: targeting, collection, and emittance reduction are paramount, as well as the bunch manipulation required to produce  $> 10^{12}$  muons per bunch without emittance degradation. The proton source needs to deliver a beam power of several MW, collection would be aided by magnetic fields common on neutron stars (though scaled back for application on earth), and the emittance requirements have inspired fascinating investigations into phase space manipulations that are finding applications in other facilities. The status was summarized in a White Paper submitted to “Snowmass 2013” [41].

#### 30.6.4. Plasma Acceleration and Other Advanced Concepts

: At the 1956 CERN Symposium, a paper by Veksler, in which he suggested acceleration of protons to the TeV scale using a bunch of electrons, anticipated current interest in plasma acceleration [42]. A half-century later this is more than a suggestion, with the demonstration, as a striking example, of electron energy doubling from 42 to 84 GeV over 85 cm at SLAC [43].

Whether plasma acceleration will find application in an HEP facility is not yet clear, given the necessity of staging and phase-locking acceleration in multiple plasma chambers. Maintaining beam quality and beam position as well as the acceleration of high-repetition bunch trains are also primary feasibility issues, addressed by active R&D. For recent discussions of parameters for a laser-plasma based electron positron collider, see, for example, relevant papers in an Advanced Accelerator Concepts Workshop [44] and the ICFA-ICUIL White Paper from 2011 [45].

Additional approaches aiming at accelerating gradients higher, or much higher, than those achievable with conventional metal cavities include the use of dielectric materials and, for the long-term future, crystals. Combining several innovative ideas, even a linear crystal muon collider driven by X-ray lasers has been proposed [46].

Not only the achievable accelerating gradient, but also the overall power efficiency, e.g. the attainable luminosity as a function of electrical input power, will determine the suitability of any novel technology for use in future high-energy accelerators.

#### References:

1. E. D. Courant and H. S. Snyder, *Ann. Phys.* **3**, 1 (1958). This is the classic article on the alternating gradient synchrotron.
2. A.W. Chao *et al.*, eds., *Handbook of Accelerator Physics and Engineering*, World Science Publishing Co. (Singapore, 2nd edition, 2013.), Sec. 2.1, 2.2.
3. H. Wiedemann, *Handbook of Accelerator Physics and Engineering*, *ibid*, Sec. 3.1.
4. C. Amsler *et al.* (Particle Data Group), *Phys. Lett.* **B667**, 1 (2008) [[http://pdg.lbl.gov/2008/reviews/contents\\_sports.html](http://pdg.lbl.gov/2008/reviews/contents_sports.html)].
5. M.A. Furman and M.S. Zisman, *Handbook of Accelerator Physics and Engineering*, *ibid*, Sec. 4.1.
6. M.A. Furman, *Handbook of Accelerator Physics and Engineering*, *ibid*, Sec. 2.4.14.
7. <http://ab-abp-rlc.web.cern.ch/ab-abp-rlc-ecloud/>. This site contains many references as well as videos of electron cloud simulations.
8. D. Möhl *et al.*, *Phys. Reports* **58**, 73 (1980).
9. M. Blaskiewicz, J.M. Brennan, and K. Mernick, *Phys. Rev. Lett.* **105**, 094801 (2010).
10. J. M. Brennan, M. Blaskiewicz, K. Mernick, “Stochastic Cooling in RHIC,” *Proc. IPAC’12, New Orleans*.
11. G.I. Budker, *Proc. Int. Symp. Electron & Positron Storage Rings*, (1966).
12. H.T. Edwards, “The Tevatron Energy Doubler: A Superconducting Accelerator,” *Ann. Rev. Nucl. and Part. Sci.* **35**, 605 (1985).

13. Brief history at [http://en.wikipedia.org/wiki/Hadron\\_Elektron\\_Ring\\_Anlage](http://en.wikipedia.org/wiki/Hadron_Elektron_Ring_Anlage).
14. R. Assmann *et al.*, Nucl. Phys. (Proc. Supp.) **B9**, 17 (2002).
15. N. Phinney, arXiv:physics/0010008 (2000).
16. L. Evans, "The Large Hadron Collider," Ann. Rev. Nucl. and Part. Sci. **61**, 435 (2011).
17. M. Draper, ed., *Proc. Chamonix 2014: LHC Performance Workshop*, CERN-2015-002, <http://cds.cern.ch/record/2020930> (2015).
18. Detailed information from the multi-volume design report to present status may be found at <http://lhc.web.cern.ch/lhc/>.
19. S. van der Meer, "Calibration of the Effective Beam Height at the ISR," CERN-ISR-PO/68-31 (1968).
20. ATLAS Collaboration, "Improved Luminosity Determination in  $pp$  Collisions at  $\sqrt{s} = 7$  TeV using the ATLAS Detector at the LHC," Eur. Phys. J. **C73**, 2518 (2013).
21. <http://aliceinfo.cern.ch/Public/Welcome.html>.
22. <http://hilumilhc.web.cern.ch>.
23. An overview of electron-positron colliders past and present may be found in ICFA Beam Dynamics Newsletter No. 46, April 2009, <http://www-bd.fnal.gov/icfabd/>. A day-by-day account of the luminosity progress at KEK-B may be found at [http://belle.kek.jp/bdocs/lumi\\_belle.png](http://belle.kek.jp/bdocs/lumi_belle.png).
24. M. Zobov *et al.*, "Test of 'Crab-Waist' Collisions at the DAΦNE  $\Phi$  Factory," Phys. Rev. Lett. **104**, 174801 (2010).
25. M. Harrison, T. Ludlam, and S. Ozaki, eds., "Special Issue: The Relativistic Heavy Ion Collider Project: RHIC and its Detectors," Nucl. Instrum. Methods **A499**, 1 (2003).
26. J.L. Abelleira *et al.*, J. Phys. **G39**, 075001 (2012).
27. J.L. Abelleira *et al.*, arXiv:1211.5102 (2012).
28. S.A. Bogacz *et al.*, arXiv:1208.2827 (2012).
29. W. Chou *et al.*, arXiv:1305.5202 (2013).
30. F. Zimmermann, "Tutorial on Linear Colliders," AIP Conf. Proc. **592**, 494 (2001).
31. [www.linearcollider.org/ILC/Publications/Technical-Design-Report](http://www.linearcollider.org/ILC/Publications/Technical-Design-Report).
32. <http://lcd.web.cern.ch/lcd/CDR/CDR.html>.
33. <http://www.cern.ch/fcc>.
34. <http://cepc.ihep.ac.cn>.
35. <http://vlhc.org>.
36. L. Rossi, E. Todesco, arXiv:1108.1619, in *Proc. HE-LHC10*, Malta, 14–16 October 2010, *CERN Yellow Report CERN-2011-003*.
37. F. Zimmermann *et al.*, "Challenges for Highest Energy Circular Colliders," *Proc. IPAC'14 Dresden*.
38. C.O. Dominguez, F. Zimmermann, "Beam Parameters and Luminosity Time Evolution for an 80-km VHE-LHC," *Proc. IPAC'13 Shanghai*.
39. <http://www.cern.ch/tlep>.
40. [http://en.wikipedia.org/wiki/Neutrino\\_Factory](http://en.wikipedia.org/wiki/Neutrino_Factory).
41. J.P. Delahaye *et al.*, arXiv:1308.0494 (2013).
42. V.I. Veksler, CERN Symposium on High Energy Accelerators and Pion Physics, 11–23 June 1956, p. 80. This paper may be downloaded from <http://cdsweb.cern.ch/record/1241563?ln=en>.
43. I. Blumenfeld *et al.*, Nature **445**, 741 (2007).
44. Advanced Accelerator Concepts, edited by R. Zgadzaj, E. Gaul and M. Downer, *AIP Conference Proceedings 1507*, Austin TX 10 – 15 June 2012.
45. White Paper of ICFA-ICUIL Joint Task Force — High Power Laser Technology for Accelerators, published in ICFA Beam Dynamics Newsletter No. 56, December 2011, <http://www-bd.fnal.gov/icfabd/>.
46. V. Shiltsev, Sov. Phys. Usp. **55**, 965 (2012).

## HIGH-ENERGY COLLIDER PARAMETERS: $e^+e^-$ Colliders (I)

Updated in January 2016 with numbers received from representatives of the colliders (contact S. Pagan Griso, LBNL). The table shows the parameter values achieved. Quantities are, where appropriate, r.m.s.; unless noted otherwise, energies refer to beam energy;  $H$  and  $V$  indicate horizontal and vertical directions; s.c. stands for superconducting. Parameters for the defunct SPEAR, DORIS, PETRA, PEP, TRISTAN, and VEPP-2M colliders may be found in our 1996 edition (Phys. Rev. **D54**, 1 July 1996, Part I).

	VEPP-2000 (Novosibirsk)	VEPP-4M (Novosibirsk)	BEPC (China)	BEPC-II (China)	DAΦNE (Frascati)
Physics start date	2010	1994	1989	2008	1999
Physics end date	—	—	2005	—	—
Maximum beam energy (GeV)	1.0	6	2.5	1.89 (2.3 max)	0.510
Delivered integrated luminosity per exp. ( $\text{fb}^{-1}$ )	0.030	0.027	0.11	10.3	$\approx 4.7$ in 2001-2007 $\approx 2.7$ w/crab-waist $\approx 1.8$ since Nov 2014
Luminosity ( $10^{30} \text{ cm}^{-2}\text{s}^{-1}$ )	100	20	12.6 at 1.843 GeV 5 at 1.55 GeV	853	453
Time between collisions ( $\mu\text{s}$ )	0.04	0.6	0.8	0.008	0.0027
Full crossing angle ( $\mu \text{ rad}$ )	0	0	0	$2.2 \times 10^4$	$5 \times 10^4$
Energy spread (units $10^{-3}$ )	0.64	1	0.58 at 2.2 GeV	0.52	0.40
Bunch length (cm)	4	5	$\approx 5$	$\approx 1.5$	low current: 1 at 15mA: 2
Beam radius ( $10^{-6} \text{ m}$ )	125 (round)	$H$ : 1000 $V$ : 30	$H$ : 890 $V$ : 37	$H$ : 358 $V$ : 4.8	$H$ : 260 $V$ : 4.8
Free space at interaction point (m)	$\pm 1$	$\pm 2$	$\pm 2.15$	$\pm 0.63$	$\pm 0.295$
Luminosity lifetime (hr)	continuous	2	7–12	1.5	0.2
Turn-around time (min)	continuous	18	32	15	2 (topping up)
Injection energy (GeV)	0.2–1.0	1.8	1.55	1.89	on energy
Transverse emittance ( $10^{-9} \text{ m}$ )	$H$ : 250 $V$ : 250	$H$ : 200 $V$ : 20	$H$ : 660 $V$ : 28	$H$ : 128 $V$ : 1.73	$H$ : 260 $V$ : 2.6
$\beta^*$ , amplitude function at interaction point (m)	$H$ : 0.06 – 0.11 $V$ : 0.06 – 0.10	$H$ : 0.75 $V$ : 0.05	$H$ : 1.2 $V$ : 0.05	$H$ : 1.0 $V$ : 0.0135	$H$ : 0.26 $V$ : 0.009
Beam-beam tune shift per crossing (units $10^{-4}$ )	$H$ : 750 $V$ : 750	500	350	390	440 (crab-waist test)
RF frequency (MHz)	172	180	199.53	499.8	356
Particles per bunch (units $10^{10}$ )	16	15	20 at 2 GeV 11 at 1.55 GeV	3.8	$e^-$ : 3.2 $e^+$ : 2.1
Bunches per ring per species	1	2	1	92	100 to 105 (120 buckets)
Average beam current per species (mA)	150	80	40 at 2 GeV 22 at 1.55 GeV	701	$e^-$ : 1250 $e^+$ : 800
Circumference or length (km)	0.024	0.366	0.2404	0.23753	0.098
Interaction regions	2	1	2	1	1
Magnetic length of dipole (m)	1.2	2	1.6	outer ring: 1.6 inner ring: 1.41	outer ring: 1.2 inner ring: 1
Length of standard cell (m)	12	7.2	6.6	outer ring: 6.6 inner ring: 6.2	n/a
Phase advance per cell (deg)	$H$ : 738 $V$ : 378	65	$\approx 60$	60–90 non-standard cells	—
Dipoles in ring	8	78	40 + 4 weak	84 + 8 weak	8
Quadrupoles in ring	20	150	68	134+2 s.c.	48
Peak magnetic field (T)	2.4	0.6	0.903 at 2.8 GeV	outer ring: 0.677 inner ring: 0.766	1.2

HIGH-ENERGY COLLIDER PARAMETERS:  $e^+e^-$  Colliders (II)

Updated in January 2016 with numbers received from representatives of the colliders (contact S. Pagan Griso, LBNL). The table shows the parameter values achieved. For future colliders, design values are quoted. Quantities are, where appropriate, r.m.s.; unless noted otherwise, energies refer to beam energy;  $H$  and  $V$  indicate horizontal and vertical directions; s.c. stands for superconducting.

	CESR (Cornell)	CESR-C (Cornell)	LEP (CERN)	SLC (SLAC)	ILC (TBD)	CLIC (TBD)
Physics start date	1979	2002	1989	1989	TBD	TBD
Physics end date	2002	2008	2000	1998	—	—
Maximum beam energy (GeV)	6	6	100 - 104.6	50	250 (upgradeable to 500)	1500 (first phase: 190)
Delivered integrated luminosity per experiment ( $\text{fb}^{-1}$ )	41.5	2.0	0.221 at Z peak 0.501 at 65 – 100 GeV 0.275 at >100 GeV	0.022	—	—
Luminosity ( $10^{30} \text{ cm}^{-2}\text{s}^{-1}$ )	1280 at 5.3 GeV	76 at 2.08 GeV	24 at Z peak 100 at > 90 GeV	2.5	$1.5 \times 10^{4\dagger}$	$6 \times 10^4$
Time between collisions ( $\mu\text{s}$ )	0.014 to 0.22	0.014 to 0.22	22	8300	$0.55^\dagger$	$0.0005^\dagger$
Full crossing angle ( $\mu \text{ rad}$ )	$\pm 2000$	$\pm 3300$	0	0	14000	20000
Energy spread (units $10^{-3}$ )	0.6 at 5.3 GeV	0.82 at 2.08 GeV	0.7→1.5	1.2	1	3.4
Bunch length (cm)	1.8	1.2	1.0	0.1	0.03	0.0044
Beam radius ( $\mu\text{m}$ )	$H$ : 460 $V$ : 4	$H$ : 340 $V$ : 6.5	$H$ : 200 → 300 $V$ : 2.5 → 8	$H$ : 1.5 $V$ : 0.5	$H$ : 0.474 $V$ : 0.0059	$H$ : 0.045 * $V$ : 0.0009
Free space at interaction point (m)	$\pm 2.2$ ( $\pm 0.6$ to REC quads)	$\pm 2.2$ ( $\pm 0.3$ to PM quads)	$\pm 3.5$	$\pm 2.8$	$\pm 3.5$	$\pm 3.5$
Luminosity lifetime (hr)	2–3	2–3	20 at Z peak 10 at > 90 GeV	—	n/a	n/a
Turn-around time (min)	5 (topping up)	1.5 (topping up)	50	120 Hz (pulsed)	n/a	n/a
Injection energy (GeV)	1.8–6	1.5–6	22	45.64	n/a	n/a
Transverse emittance ( $10^{-9} \text{ m}$ )	$H$ : 210 $V$ : 1	$H$ : 120 $V$ : 3.5	$H$ : 20–45 $V$ : 0.25 → 1	$H$ : 0.5 $V$ : 0.05	$H$ : 0.02 $V$ : $7 \times 10^{-5}$	$H$ : $2.2 \times 10^{-4}$ $V$ : $6.8 \times 10^{-6}$
$\beta^*$ , amplitude function at interaction point (m)	$H$ : 1.0 $V$ : 0.018	$H$ : 0.94 $V$ : 0.012	$H$ : 1.5 $V$ : 0.05	$H$ : 0.0025 $V$ : 0.0015	$H$ : 0.01 $V$ : $5 \times 10^{-4}$	$H$ : 0.0069 $V$ : $6.8 \times 10^{-5}$
Beam-beam tune shift per crossing ( $10^{-4}$ ) or disruption	$H$ : 250 $V$ : 620	$e^-$ : 420 ( $H$ ), 280 ( $V$ ) $e^+$ : 410 ( $H$ ), 270 ( $V$ )	830	0.75 ( $H$ ) 2.0 ( $V$ )	n/a	7.7
RF frequency (MHz)	500	500	352.2	2856	1300	11994
Particles per bunch (units $10^{10}$ )	1.15	4.7	45 in collision 60 in single beam	4.0	2	0.37
Bunches per ring per species	9 trains of 5 bunches	8 trains of 3 bunches	4 trains of 1 or 2	1	1312	312 (in train)
Average beam current per species (mA)	340	72	4 at Z peak 4→6 at > 90 GeV	0.0008	6 (in pulse)	1205 (in train)
Beam polarization (%)	—	—	55 at 45 GeV 5 at 61 GeV	$e^-$ : 80	$e^-$ : > 80% $e^+$ : < 60%	$e^-$ : 70% at IP
Circumference or length (km)	0.768	0.768	26.66	1.45 +1.47	31	50
Interaction regions	1	1	4	1	1	1
Magnetic length of dipole (m)	1.6–6.6	1.6–6.6	11.66/pair	2.5	n/a	n/a
Length of standard cell (m)	16	16	79	5.2	n/a	n/a
Phase advance per cell (deg)	45–90 (no standard cell)	45–90 (no standard cell)	102/90	108	n/a	n/a
Dipoles in ring	86	84	3280 + 24 inj. + 64 weak	460+440	n/a	n/a
Quadrupoles in ring	101 + 4 s.c.	101 + 4 s.c.	520 + 288 + 8 s.c.	—	n/a	n/a
Peak magnetic field (T)	0.3 / 0.8 at 8 GeV	0.3 / 0.8 at 8 GeV, 2.1 wigglers at 1.9 GeV	0.135	0.597	n/a	n/a

<sup>†</sup>Time between bunch trains: 200ms (ILC) and 20ms (CLIC).

<sup>‡</sup>Geometrical luminosity. The actual value may vary by  $\approx 20\%$  depending on assumptions.

\*Effective beam size including non-linear and chromatic effects.

## HIGH-ENERGY COLLIDER PARAMETERS: $e^+e^-$ Colliders (III)

Updated in January 2016 with numbers received from representatives of the colliders (contact S. Pagan Griso, LBNL). The table shows the parameter values achieved. For future colliders, design values are quoted. Quantities are, where appropriate, r.m.s.; unless noted otherwise, energies refer to beam energy;  $H$  and  $V$  indicate horizontal and vertical directions; s.c. stands for superconducting.

	KEKB (KEK)	PEP-II (SLAC)	SuperKEKB (KEK)
Physics start date	1999	1999	2017
Physics end date	2010	2008	—
Maximum beam energy (GeV)	$e^-$ : 8.33 (8.0 nominal) $e^+$ : 3.64 (3.5 nominal)	$e^-$ : 7–12 (9.0 nominal) $e^+$ : 2.5–4 (3.1 nominal)	$e^-$ : 7 $e^+$ : 4
Delivered integrated luminosity per exp. ( $\text{fb}^{-1}$ )	1040	557	—
Luminosity ( $10^{30} \text{ cm}^{-2}\text{s}^{-1}$ )	21083	12069 (design: 3000)	$8 \times 10^5$
Time between collisions ( $\mu\text{s}$ )	0.00590 or 0.00786	0.0042	0.004
Full crossing angle ( $\mu$ rad)	$\pm 11000^\dagger$	0	$\pm 41500$
Energy spread (units $10^{-3}$ )	0.7	$e^-/e^+$ : 0.61/0.77	$e^-/e^+$ : 0.64/0.81
Bunch length (cm)	0.65	$e^-/e^+$ : 1.1/1.0	$e^-/e^+$ : 0.5/0.6
Beam radius ( $\mu\text{m}$ )	H: 124 ( $e^-$ ), 117 ( $e^+$ ) V: 1.9	H: 157 V: 4.7	$e^-$ : 11 (H), 0.062 (V) $e^+$ : 10 (H), 0.048 (V)
Free space at interaction point (m)	+0.75/−0.58 (+300/−500) mrad cone	$\pm 0.2$ , $\pm 300$ mrad cone	$e^-$ : +1.20/−1.28, $e^+$ : +0.78/−0.73 (+300/−500) mrad cone
Luminosity lifetime (hr)	continuous	continuous	continuous
Turn-around time (min)	continuous	continuous	continuous
Injection energy (GeV)	$e^-/e^+$ : 8.0/3.5 (nominal)	$e^-/e^+$ : 9.0/3.1 (nominal)	$e^-/e^+$ : 7/4
Transverse emittance ( $10^{-9}$ m)	$e^-$ : 24 (57*) (H), 0.61 (V) $e^+$ : 18 (55*) (H), 0.56 (V)	$e^-$ : 48 (H), 1.8 (V) $e^+$ : 24 (H), 1.8 (V)	$e^-$ : 4.6 (H), 0.013 (V) $e^+$ : 3.2 (H), 0.0086 (V)
$\beta^*$ , amplitude function at interaction point (m)	$e^-$ : 1.2 (0.27*) (H), 0.0059 (V) $e^+$ : 1.2 (0.23*) (H), 0.0059 (V)	$e^-$ : 0.50 (H), 0.012 (V) $e^+$ : 0.50 (H), 0.012 (V)	$e^-$ : 0.025 (H), $3 \times 10^{-4}$ (V) $e^+$ : 0.032 (H), $2.7 \times 10^{-4}$ (V)
Beam-beam tune shift per crossing (units $10^{-4}$ )	$e^-$ : 1020 (H), 900 (V) $e^+$ : 1270 (H), 1290 (V)	$e^-$ : 703 (H), 498 (V) $e^+$ : 510 (H), 727 (V)	$e^-$ : 12 (H), 807 (V) $e^+$ : 28 (H), 881 (V)
RF frequency (MHz)	508.887	476	508.887
Particles per bunch (units $10^{10}$ )	$e^-/e^+$ : 4.7/6.4	$e^-/e^+$ : 5.2/8.0	$e^-/e^+$ : 6.53/9.04
Bunches per ring per species	1585	1732	2500
Average beam current per species (mA)	$e^-/e^+$ : 1188/1637	$e^-/e^+$ : 1960/3026	$e^-/e^+$ : 2600/3600
Beam polarization (%)	—	—	—
Circumference or length (km)	3.016	2.2	3.016
Interaction regions	1	1	1
Magnetic length of dipole (m)	$e^-/e^+$ : 5.86/0.915	$e^-/e^+$ : 5.4/0.45	$e^-/e^+$ : 5.9/4.0
Length of standard cell (m)	$e^-/e^+$ : 75.7/76.1	15.2	$e^-/e^+$ : 75.7/76.1
Phase advance per cell (deg)	450	$e^-/e^+$ : 60/90	450
Dipoles in ring	$e^-/e^+$ : 116/112	$e^-/e^+$ : 192/192	$e^-/e^+$ : 116/112
Quadrupoles in ring	$e^-/e^+$ : 452/452	$e^-/e^+$ : 290/326	$e^-/e^+$ : 466/460
Peak magnetic field (T)	$e^-/e^+$ : 0.25/0.72	$e^-/e^+$ : 0.18/0.75	$e^-/e^+$ : 0.22/0.19

$^\dagger$ KEKB was operated with crab crossing from 2007 to 2010.

\*With dynamic beam-beam effect.

HIGH-ENERGY COLLIDER PARAMETERS:  $e\bar{p}$ ,  $p\bar{p}$ ,  $pp$  Colliders

Updated in January 2016 with numbers received from representatives of the colliders (contact S. Pagan Griso, LBNL). The table shows the parameter values achieved. For LHC, the parameters expected at the ATLAS and CMS experiments for a high-luminosity upgrade (HL-LHC) are also given. Parameters for the defunct  $Sp\bar{p}S$  collider may be found in our 2002 edition (Phys. Rev. **D66**, 010001 (2002)). Quantities are, where appropriate, r.m.s.; unless noted otherwise, energies refer to beam energy;  $H$  and  $V$  indicate horizontal and vertical directions; s.c. stands for superconducting.

	HERA (DESY)	TEVATRON* (Fermilab)	RHIC (Brookhaven)	LHC (CERN)		
Physics start date	1992	1987	2001	2009	2015	2024 (HL-LHC)
Physics end date	2007	2011	—	—		
Particles collided	$e\bar{p}$	$p\bar{p}$	$pp$ (polarized)	$pp$		
Maximum beam energy (TeV)	$e$ : 0.030 $p$ : 0.92	0.980	0.255 53% polarization	4.0	6.5	7.0
Maximum delivered integrated luminosity per exp. ( $\text{fb}^{-1}$ )	0.8	12	0.38 at 100 GeV 0.75 at 250/255 GeV	23.3 at 4.0 TeV 6.1 at 3.5 TeV	4.2	250/y
Luminosity ( $10^{30} \text{ cm}^{-2}\text{s}^{-1}$ )	75	431	245 (pk) 160 (avg)	$7.7 \times 10^3$	$5 \times 10^3$	$5.0 \times 10^4$ (leveled)
Time between collisions (ns)	96	396	107	49.90	24.95	24.95
Full crossing angle ( $\mu\text{ rad}$ )	0	0	0	290	290	590
Energy spread (units $10^{-3}$ )	$e$ : 0.91 $p$ : 0.2	0.14	0.15	0.1445	0.105	0.123
Bunch length (cm)	$e$ : 0.83 $p$ : 8.5	$p$ : 50 $\bar{p}$ : 45	60	9.4	9	9
Beam radius ( $10^{-6} \text{ m}$ )	$e$ : 110( $H$ ), 30( $V$ ) $p$ : 111( $H$ ), 30( $V$ )	$p$ : 28 $\bar{p}$ : 16	85	18.8	21	7
Free space at interaction point (m)	$\pm 2$	$\pm 6.5$	16	38	38	38
Initial luminosity decay time, $-L/(dL/dt)$ (hr)	10	6 (avg)	7.5	$\approx 6$	$\approx 30$	$\approx 6$ (leveled)
Turn-around time (min)	$e$ : 75, $p$ : 135	90	25	180	134	180
Injection energy (TeV)	$e$ : 0.012 $p$ : 0.040	0.15	0.023	0.450	0.450	0.450
Transverse emittance ( $10^{-9} \text{ m}$ )	$e$ : 20( $H$ ), 3.5( $V$ ) $p$ : 5( $H$ ), 5( $V$ )	$p$ : 3 $\bar{p}$ : 1	13	0.59	0.5	0.34
$\beta^*$ , ampl. function at interaction point (m)	$e$ : 0.6( $H$ ), 0.26( $V$ ) $p$ : 2.45( $H$ ), 0.18( $V$ )	0.28	0.65	0.6	0.8	0.15
Beam-beam tune shift per crossing (units $10^{-4}$ )	$e$ : 190( $H$ ), 450( $V$ ) $p$ : 12( $H$ ), 9( $V$ )	$p$ : 120 $\bar{p}$ : 120	73	72	37	110
RF frequency (MHz)	$e$ : 499.7 $p$ : 208.2/52.05	53	accel: 9 store: 28	400.8	400.8	400.8
Particles per bunch (units $10^{10}$ )	$e$ : 3 $p$ : 7	$p$ : 26 $\bar{p}$ : 9	18.5	16	12	22
Bunches per ring per species	$e$ : 189 $p$ : 180	36	111	1380	2244 2232 (i.r. $1/5^\dagger$ )	2748 2736 (i.r. $1/5^\dagger$ )
Average beam current per species (mA)	$e$ : 40 $p$ : 90	$p$ : 70 $\bar{p}$ : 24	257	400	467	1200
Circumference (km)	6.336	6.28	3.834	26.659		
Interaction regions	2 colliding beams 1 fixed target ( $e$ beam)	2 high $\mathcal{L}$	6 total, 2 high $\mathcal{L}$	4 total, 2 high $\mathcal{L}$		
Magnetic length of dipole (m)	$e$ : 9.185 $p$ : 8.82	6.12	9.45	14.3		
Length of standard cell (m)	$e$ : 23.5 $p$ : 47	59.5	29.7	106.90		
Phase advance per cell (deg)	$e$ : 60 $p$ : 90	67.8	84	90		
Dipoles in ring	$e$ : 396 $p$ : 416	774	192 per ring + 12 common	1232 main dipoles		
Quadrupoles in ring	$e$ : 580 $p$ : 280	216	246 per ring	482 2-in-1 24 1-in-1		
Magnet types	$e$ : C-shaped $p$ : s.c., collared, warm iron	s.c., $\cos\theta$ warm iron	s.c., $\cos\theta$ cold iron	s.c., 2 in 1 cold iron		
Peak magnetic field (T)	$e$ : 0.274, $p$ : 5	4.4	3.5	8.3 <sup>‡</sup>		

\*Additional TEVATRON parameters:  $\bar{p}$  source accum. rate:  $25 \times 10^{10} \text{ hr}^{-1}$ ; max. no. of  $\bar{p}$  stored:  $3.4 \times 10^{12}$  (Accumulator),  $6.1 \times 10^{12}$  (Recycler).

<sup>†</sup>Number of bunches colliding at the interaction regions (i.r.) 1 (ATLAS) and 5 (CMS).

<sup>‡</sup>Value for design beam energy of 7 TeV.

## HIGH-ENERGY COLLIDER PARAMETERS: Heavy Ion Colliders

Updated in January 2016 with numbers received from representatives of the colliders (contact S. Pagan Griso, LBNL). The table shows the parameter values achieved. For LHC, the parameters expected at the ATLAS experiment for running in 2016 and the design values for a high-luminosity upgrade are also given. Quantities are, where appropriate, r.m.s.; unless noted otherwise, energies refer to beam energy; s.c. stands for superconducting. pk and avg denote peak and average values.

	RHIC (Brookhaven)		LHC (CERN)			
Physics start date	2000	2012 / 2012 / 2004 / 2014 2002 / 2015 / 2015	2010	2012	2016 (expected)	$\geq 2021$ (high lum.) <sup>‡</sup>
Physics end date	—		—			
Particles collided	Au Au	U U / Cu Au / Cu Cu / h Au d Au / p Au / p Al	Pb Pb	p Pb	p Pb	Pb Pb
Maximum beam energy (TeV/n)	0.1	0.1	2.51	p: 4 Pb: 1.58	p: 6.5 Pb: 2.56	2.76
$\sqrt{s_{NN}}$ (TeV)	0.2	0.2	5.02	5.0	8.16	5.5
Max. delivered int. nucleon-pair lumin. per exp. (pb <sup>-1</sup> )	1484 (at 100 GeV/n)	21 / 167 / 65 / 43 103 / 125 / 64 (all at 100 GeV/n)	30.3	6.6	$\approx 10/y$	$\approx 75 - 90/y$
Luminosity (10 <sup>27</sup> cm <sup>-2</sup> s <sup>-1</sup> )	pk: 8.4 avg: 8.0	pk: 0.9 / 12 / 20 / 170 270 / 880 / 7150 avg: 0.6 / 10 / 0.8 / 100 140 / 450 / 4000	3.6	100 (leveled) 116 (ATLAS/CMS)	$\approx 500$	6 (leveled)
Time between collisions (ns)	107	107 / 107 / 321 / 107 107 / 107 / 107	99.8 / 149.7	199.6 / 224.6	99.8 / 149.7	49.9
Full crossing angle ( $\mu$ rad)	0	0	290	120	290	> 200
Energy spread (units 10 <sup>-3</sup> )	0.75	0.75	0.11	0.11	0.11	0.11
Bunch length (cm)	30	30	8.0	p / Pb: 9 / 11.5	p / Pb: 9 / 11.5	7.9
Beam radius (10 <sup>-6</sup> m)	55	50 / 160 / 145 / 135 145 / 145 / 145	55	p: 19 Pb: 27	17	16
Free space at interaction point (m)	16	16	38	38	38	38
Initial luminosity decay time, $-L/(dL/dt)$ (hr)	1	-0.35 <sup>†</sup> / $\infty^{\dagger}$ / 1.8 / 0.6 1.5 / 0.5 / 0.25	2.6	$\approx 6$	$\approx 2$	$\approx 2$
Turn-around time (min)	30	60 / 160 / 90 / 45 90 / 60 / 50	$\approx 180$	$\approx 240$	$\approx 180$	$\approx 180$
Injection energy (TeV/n)	0.011	0.011	0.177	p / Pb: 0.45 / 0.177	p / Pb: 0.45 / 0.177	0.177
Transverse emittance (10 <sup>-9</sup> m)	6	4 / 11 / 23 / 18 25 / 25 / 23	1.5	p: 0.5 Pb: 0.9	0.29	0.5
$\beta^*$ , ampl. function at interaction point (m)	0.5	0.7 / 0.7 / 0.9 / 1.0 0.85 / 0.8 / 0.8	0.8	0.8	0.5	0.5
Beam-beam tune shift per crossing (units 10 <sup>-4</sup> )	25	7 / 14 (Cu), 14 (Au) / 30 42 (h), 22 (Au) / 21 (d), 17 (Au) 53 (p), 41 (Au) / 73 (p) 57 (Au)	9	p: 9 Pb: 10	10	10
RF frequency (MHz)	accel: 28 store: 197	accel: 28 store: 197	400.8	400.8	400.8	400.8
Particles per bunch (units 10 <sup>10</sup> )	0.16	0.03 / 0.4 (Cu), 0.13 (Au) / 0.45 4.5 (h), 0.13 (Au) / 10 (d), 0.1 (Au) 22.5 (p), 0.16 (Au) / 24 (p), 1.1 (Al)	0.019 (r.m.s.)	p: 1.6 Pb: 0.014	p: 1.8 Pb: 0.019	0.017
Bunches per ring per species	111	111 / 111 / 37 / 111 95 / 111 / 111	518	338	518	$\approx 1100$
Average beam current per species (mA)	176	38 / 159 (Cu), 138 (Au) / 60 125 (h), 143 (Au) / 119 (d), 94 (Au) 312 (p), 176 (Au) / 333 (p), 199 (Al)	14.9	p: 9.7 Pb: 7	p: 17 Pb: 15	28
Circumference (km)	3.834		26.659			
Interaction regions	6 total, 2 high $\mathcal{L}$		3 high $\mathcal{L}$ + 1			
Magnetic length of dipole (m)	9.45		14.3			
Length of standard cell (m)	29.7		106.90			
Phase advance per cell (deg)	93	84 / 84 / 84 / 93 84 (d), 93 (Au) / 84 (p), 93 (Au) 84 (p), 93 (Al)	90			
Dipoles in ring	192 per ring, + 12 common		1232, main dipoles			
Quadrupoles in ring	246 per ring		482 2-in-1, 24 1-in-1			
Magnet Type	s.c. $\cos\theta$ , cold iron		s.c., 2 in 1, cold iron			
Peak magnetic field (T)	3.5		8.3			

<sup>†</sup>Negative or infinite decay time is effect of cooling.

<sup>‡</sup>High luminosity upgrade expected  $\geq 2021$ ; will extend throughout HL-LHC running. Very preliminary, conservative estimates.

## 32. NEUTRINO BEAM LINES AT HIGH-ENERGY PROTON SYNCHROTRONS

Revised January 2016 with numbers verified by representatives of the synchrotrons (contact C.-J. Lin, LBNL). For existing (future) neutrino beam lines the latest achieved (design) values are given.

The main source of neutrinos at proton synchrotrons is from the decay of pions and kaons produced by protons striking a nuclear target. There are different schemes to focus the secondary particles to enhance neutrino flux and/or tune the neutrino energy profile. In wide-band beams (WBB), the neutrino parent mesons are focused over a wide momentum range to obtain maximum neutrino intensity. In narrow-band beams (NBB), the secondary particles are first momentum-selected to produce a monochromatic parent beam. Another approach to generate a narrow-band neutrino spectrum is to select neutrinos that are emitted off-axis relative to the momentum of the parent mesons. For a comprehensive review of the topic, including other historical neutrino beam lines, see the article by S. E. Kopp, "Accelerator-based neutrino beams," Phys. Rept. **439**, 101 (2007).

	PS (CERN)				SPS (CERN)				PS (KEK)	Main Ring (JPARC)
Date	1963	1969	1972	1983	1977	1977	1995	2006	1999	2009
Proton Kinetic Energy (GeV)	20.6	20.6	26	19	350	350	450	400	12	30 (50)
Protons per Cycle ( $10^{12}$ )	0.7	0.6	5	5	10	10	36	48	6	200 (330)
Cycle Time (s)	3	2.3	-	-	-	-	14.4	6	2.2	2.48 (3.5)
Beam Power (kW)	0.8	0.9	-	-	-	-	180	510	5	390 (750)
Target	-	-	-	-	-	-	Be	Graphite	Al	Graphite
Target Length (cm)	-	-	-	-	-	-	290	1000	66	91
Secondary Focussing	1-horn WBB	3-horn WBB	2-horn WBB	bare target	dichromatic NBB	2-horn WBB	2-horn WBB	2-horn WBB	2-horn WBB	3-horn off-axis
Decay Pipe Length (m)	-	-	-	-	-	-	110	130	200	96
$\langle E_\nu \rangle$ (GeV)	1.5	1.5	1.5	1	50,150 <sup>†</sup>	20	24.3	17	1.3	0.6
Experiments	HLBC, Spark Ch.	HLBC, Spark Ch.	GGM, Aachen-Padova	CDHS, CHARM	CDHS, CHARM, BEBC	GGM, CDHS, CHARM, BEBC	NOMAD, CHORUS	OPERA, ICARUS	K2K	T2K

	Main Ring (Fermilab)							Booster (Fermilab)	Main Injector (Fermilab)	
Date	1975	1975	1974	1979	1976	1991	1998	2002	2005	2016
Proton Kinetic Energy (GeV)	300,400	300,400	300	400	350	800	800	8	120	120
Protons per Cycle ( $10^{12}$ )	10	10	10	10	13	10	12	4.5	37	43 (49)
Cycle Time (s)	-	-	-	-	-	60	60	0.2	2	1.333
Beam Power (kW)	-	-	-	-	-	20	25	29	350	580 (700)
Target	-	-	-	-	-	-	BeO	Be	Graphite	Graphite
Target Length (cm)	-	-	-	-	-	-	31	71	95	120
Secondary Focussing	bare target	quad trip., SSBT	dichromatic NBB	2-horn WBB	1-horn WBB	quad trip.	SSQT WBB	1-horn WBB	2-horn WBB	2-horn off-axis
Decay Pipe Length (m)	350	350	400	400	400	400	400	50	675	675
$\langle E_\nu \rangle$ (GeV)	40	50,180 <sup>†</sup>	50,180 <sup>†</sup>	25	100	90,260	70,180	1	3-20 <sup>‡</sup>	2
Experiments	HPWF	CITF, HPWF	CITF, HPWF, 15' BC	15' BC	HPWF 15' BC	15' BC, CCFRR	NuTeV	MiniBooNE, SciBooNE, MicroBooNE	MINOS, MINERνA	NOνA, MINERνA, MINOS+

<sup>†</sup>Pion and kaon peaks in the momentum-selected channel.

<sup>‡</sup>Tunable WBB energy spectrum.



## 33. PASSAGE OF PARTICLES THROUGH MATTER

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Revised August 2015 by H. Bichsel (University of Washington), D.E. Groom (LBNL), and S.R. Klein (LBNL).

This review covers the interactions of photons and electrically charged particles in matter, concentrating on energies of interest for high-energy physics and astrophysics and processes of interest for particle detectors (ionization, Cherenkov radiation, transition radiation). Much of the focus is on particles heavier than electrons ( $\pi^\pm$ ,  $p$ , etc.). Although the charge number  $z$  of the projectile is included in the equations, only  $z = 1$  is discussed in detail. Muon radiative losses are discussed, as are photon/electron interactions at high to ultrahigh energies. Neutrons are not discussed.

## 33.1. Notation

The notation and important numerical values are shown in Table 33.1.

**Table 33.1:** Summary of variables used in this section. The kinematic variables  $\beta$  and  $\gamma$  have their usual relativistic meanings.

Symbol	Definition	Value or (usual) units
$m_e c^2$	electron mass $\times c^2$	0.510 998 928(11) MeV
$r_e$	classical electron radius $e^2/4\pi\epsilon_0 m_e c^2$	2.817 940 3267(27) fm
$\alpha$	fine structure constant $e^2/4\pi\epsilon_0 \hbar c$	1/137.035 999 074(44)
$N_A$	Avogadro's number	$6.022\,141\,29(27) \times 10^{23} \text{ mol}^{-1}$
$\rho$	density	$\text{g cm}^{-3}$
$x$	mass per unit area	$\text{g cm}^{-2}$
$M$	incident particle mass	MeV/ $c^2$
$E$	incident part. energy $\gamma M c^2$	MeV
$T$	kinetic energy, $(\gamma - 1) M c^2$	MeV
$W$	energy transfer to an electron in a single collision	MeV
$k$	bremsstrahlung photon energy	MeV
$z$	charge number of incident particle	
$Z$	atomic number of absorber	
$A$	atomic mass of absorber	$\text{g mol}^{-1}$
$K$	$4\pi N_A r_e^2 m_e c^2$	0.307 075 MeV $\text{mol}^{-1} \text{ cm}^2$
$I$	mean excitation energy	eV ( <i>Nota bene!</i> )
$\delta(\beta\gamma)$	density effect correction to ionization energy loss	
$\hbar\omega_p$	plasma energy $\sqrt{4\pi N_e r_e^3} m_e c^2 / \alpha$	$\sqrt{\rho \langle Z/A \rangle} \times 28.816 \text{ eV}$ $\downarrow \rho \text{ in g cm}^{-3}$
$N_e$	electron density	(units of $r_e$ ) $^{-3}$
$w_j$	weight fraction of the $j$ th element in a compound or mixture	
$n_j$	$\propto$ number of $j$ th kind of atoms in a compound or mixture	
$X_0$	radiation length	$\text{g cm}^{-2}$
$E_c$	critical energy for electrons	MeV
$E_{\mu c}$	critical energy for muons	GeV
$E_s$	scale energy $\sqrt{4\pi/\alpha} m_e c^2$	21.2052 MeV
$R_M$	Molière radius	$\text{g cm}^{-2}$

## 33.2. Electronic energy loss by heavy particles [1–33]

## 33.2.1. Moments and cross sections :

The electronic interactions of fast charged particles with speed  $v = \beta c$  occur in *single collisions with energy losses*  $W$  [1], leading to ionization, atomic, or collective excitation. Most frequently the energy losses are small (for 90% of all collisions the energy losses are less than 100 eV). In thin absorbers few collisions will take place and the total energy loss will show a large variance [1]; also see Sec. 33.2.9 below. For particles with charge  $ze$  more massive than electrons (“heavy” particles), scattering from free electrons is adequately described by the Rutherford differential cross section [2],

$$\frac{d\sigma_R(W; \beta)}{dW} = \frac{2\pi r_e^2 m_e c^2 z^2}{\beta^2} \frac{(1 - \beta^2 W/W_{\max})}{W^2}, \quad (33.1)$$

where  $W_{\max}$  is the maximum energy transfer possible in a single collision. But in matter electrons are not free.  $W$  must be finite and depends on atomic and bulk structure. For electrons bound in atoms Bethe [3] used “Born Theorie” to obtain the differential cross section

$$\frac{d\sigma_B(W; \beta)}{dW} = \frac{d\sigma_R(W; \beta)}{dW} B(W). \quad (33.2)$$

Electronic binding is accounted for by the correction factor  $B(W)$ . Examples of  $B(W)$  and  $d\sigma_B/dW$  can be seen in Figs. 5 and 6 of Ref. 1.

Bethe’s theory extends only to some energy above which atomic effects are not important. The free-electron cross section (Eq. (33.1)) can be used to extend the cross section to  $W_{\max}$ . At high energies  $\sigma_B$  is further modified by polarization of the medium, and this “density effect,” discussed in Sec. 33.2.5, must also be included. Less important corrections are discussed below.

The mean number of collisions with energy loss between  $W$  and  $W + dW$  occurring in a distance  $\delta x$  is  $N_e \delta x (d\sigma/dW) dW$ , where  $d\sigma(W; \beta)/dW$  contains all contributions. It is convenient to define the moments

$$M_j(\beta) = N_e \delta x \int W^j \frac{d\sigma(W; \beta)}{dW} dW, \quad (33.3)$$

so that  $M_0$  is the mean number of collisions in  $\delta x$ ,  $M_1$  is the mean energy loss in  $\delta x$ ,  $(M_2 - M_1)^2$  is the variance, *etc.* The number of collisions is Poisson-distributed with mean  $M_0$ .  $N_e$  is either measured in electrons/g ( $N_e = N_A Z/A$ ) or electrons/cm<sup>3</sup> ( $N_e = N_A \rho Z/A$ ). The former is used throughout this chapter, since quantities of interest ( $dE/dx$ ,  $X_0$ , *etc.*) vary smoothly with composition when there is no density dependence.

effects significantly modify the cross sections. This problem has been investigated by J.D. Jackson [9], who concluded that for hadrons (but not for large nuclei) corrections to  $dE/dx$  are negligible below energies where radiative effects dominate. While the cross section for rare hard collisions is modified, the average stopping power, dominated by many softer collisions, is almost unchanged.

### 33.2.3. Stopping power at intermediate energies :

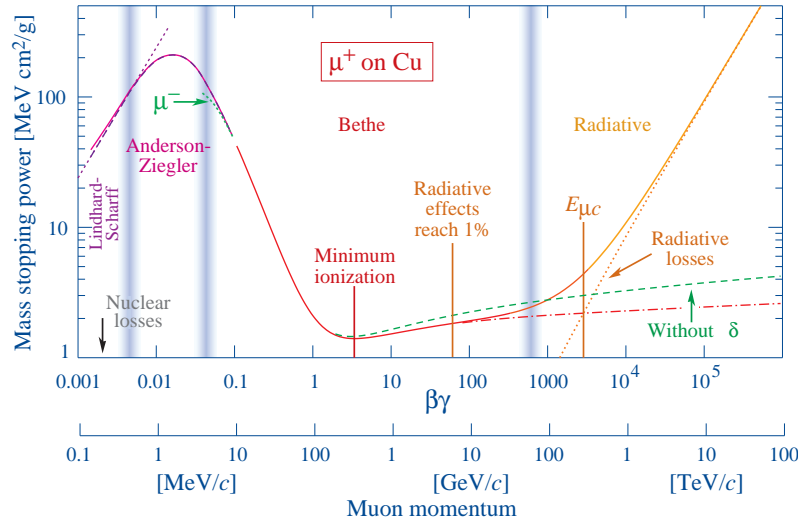
The mean rate of energy loss by moderately relativistic charged heavy particles,  $M_1/\delta x$ , is well-described by the “Bethe equation,”

$$\left\langle -\frac{dE}{dx} \right\rangle = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 W_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]. \quad (33.5)$$

It describes the mean rate of energy loss in the region  $0.1 \lesssim \beta\gamma \lesssim 1000$  for intermediate- $Z$  materials with an accuracy of a few percent.

This is the *mass stopping power*; with the symbol definitions and values given in Table 33.1, the units are MeV g<sup>-1</sup>cm<sup>2</sup>. As can be seen from Fig. 33.2,  $\langle -dE/dx \rangle$  defined in this way is about the same for most materials, decreasing slowly with  $Z$ . The *linear stopping power*, in MeV/cm, is  $\langle -dE/dx \rangle \rho$ , where  $\rho$  is the density in g/cm<sup>3</sup>.

$W_{\max}$  is defined in Sec. 33.2.2. At the lower limit the projectile velocity becomes comparable to atomic electron “velocities” (Sec. 33.2.6), and at the upper limit radiative effects begin to be important (Sec. 33.6). Both limits are  $Z$  dependent. A minor dependence on  $M$  at the highest energies is introduced through  $W_{\max}$ , but for all practical purposes  $\langle dE/dx \rangle$  in a given material is a function of  $\beta$  alone.



**Figure 33.1:** Mass stopping power ( $= \langle -dE/dx \rangle$ ) for positive muons in copper as a function of  $\beta\gamma = p/Mc$  over nine orders of magnitude in momentum (12 orders of magnitude in kinetic energy). Solid curves indicate the total stopping power. Data below the break at  $\beta\gamma \approx 0.1$  are taken from ICRU 49 [4], and data at higher energies are from Ref. 5. Vertical bands indicate boundaries between different approximations discussed in the text. The short dotted lines labeled “ $\mu^-$ ” illustrate the “Barkas effect,” the dependence of stopping power on projectile charge at very low energies [6].  $dE/dx$  in the radiative region is not simply a function of  $\beta$ .

### 33.2.2. Maximum energy transfer in a single collision :

For a particle with mass  $M$ ,

$$W_{\max} = \frac{2m_e c^2 \beta^2 \gamma^2}{1 + 2\gamma m_e/M + (m_e/M)^2}. \quad (33.4)$$

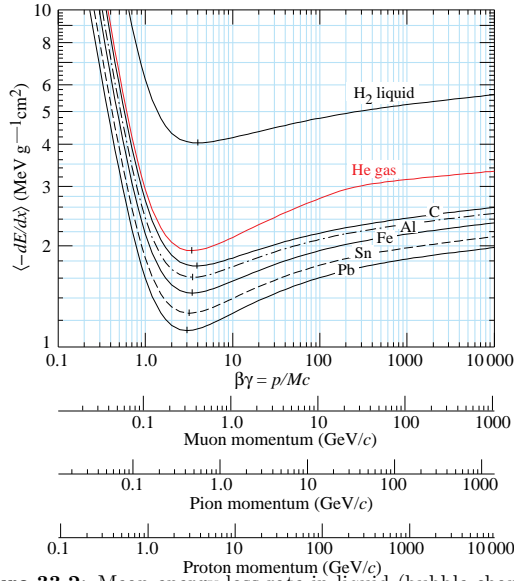
In older references [2,8] the “low-energy” approximation  $W_{\max} = 2m_e c^2 \beta^2 \gamma^2$ , valid for  $2\gamma m_e \ll M$ , is often implicit. For a pion in copper, the error thus introduced into  $dE/dx$  is greater than 6% at 100 GeV. For  $2\gamma m_e \gg M$ ,  $W_{\max} = M c^2 \beta^2 \gamma$ .

At energies of order 100 GeV, the maximum 4-momentum transfer to the electron can exceed 1 GeV/c, where hadronic structure

Few concepts in high-energy physics are as misused as  $\langle dE/dx \rangle$ . The main problem is that the mean is weighted by very rare events with large single-collision energy deposits. Even with samples of hundreds of events a dependable value for the mean energy loss cannot be obtained. Far better and more easily measured is the most probable energy loss in a detector is considerably below the mean given by the Bethe equation.

In a TPC (Sec. 34.6.5), the mean of 50%–70% of the samples with the smallest signals is often used as an estimator.

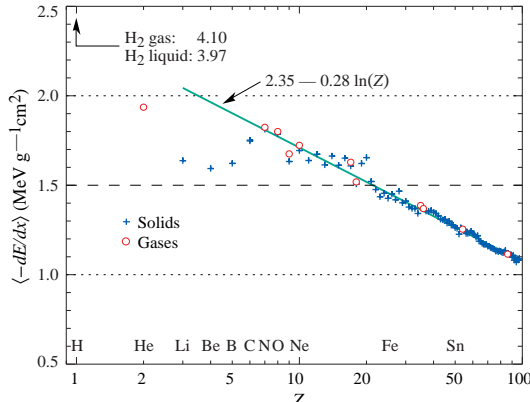
Although it must be used with cautions and caveats,  $\langle dE/dx \rangle$



**Figure 33.2:** Mean energy loss rate in liquid (bubble chamber) hydrogen, gaseous helium, carbon, aluminum, iron, tin, and lead. Radiative effects, relevant for muons and pions, are not included. These become significant for muons in iron for  $\beta\gamma \gtrsim 1000$ , and at lower momenta for muons in higher- $Z$  absorbers. See Fig. 33.23.

as described in Eq. (33.5) still forms the basis of much of our understanding of energy loss by charged particles. Extensive tables are available[4,5, [pdg.lbl.gov/AtomicNuclearProperties/](http://pdg.lbl.gov/AtomicNuclearProperties/)].

For heavy projectiles, like ions, additional terms are required to account for higher-order photon coupling to the target, and to account for the finite size of the target radius. These can change  $dE/dx$  by a factor of two or more for the heaviest nuclei in certain kinematic regimes [7].

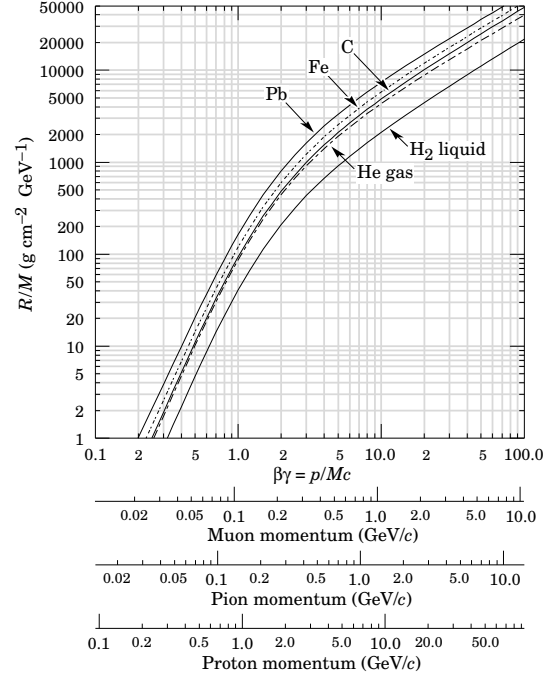


**Figure 33.3:** Mass stopping power at minimum ionization for the chemical elements. The straight line is fitted for  $Z > 6$ . A simple functional dependence on  $Z$  is not to be expected, since  $(-dE/dx)$  also depends on other variables.

The function as computed for muons on copper is shown as the “Bethe” region of Fig. 33.1. Mean energy loss behavior below this region is discussed in Sec. 33.2.6, and the radiative effects at high energy are discussed in Sec. 33.6. Only in the Bethe region is it a function of  $\beta$  alone; the mass dependence is more complicated elsewhere. The stopping power in several other materials is shown in Fig. 33.2. Except in hydrogen, particles with the same velocity have similar rates of energy loss in different materials, although there is a slow decrease in the rate of energy loss with increasing  $Z$ . The qualitative behavior difference at high energies between a gas (He in the figure) and the other materials shown in the figure is due to the density-effect correction,  $\delta(\beta\gamma)$ , discussed in Sec. 33.2.5. The stopping power functions are characterized by broad minima whose position

drops from  $\beta\gamma = 3.5$  to 3.0 as  $Z$  goes from 7 to 100. The values of minimum ionization as a function of atomic number are shown in Fig. 33.3.

In practical cases, most relativistic particles (*e.g.*, cosmic-ray muons) have mean energy loss rates close to the minimum; they are “minimum-ionizing particles,” or mip’s.



**Figure 33.4:** Range of heavy charged particles in liquid (bubble chamber) hydrogen, helium gas, carbon, iron, and lead. For example: For a  $K^+$  whose momentum is 700 MeV/c,  $\beta\gamma = 1.42$ . For lead we read  $R/M \approx 396$ , and so the range is 195 g cm<sup>-2</sup> (17 cm).

Eq. (33.5) may be integrated to find the total (or partial) “continuous slowing-down approximation” (CSDA) range  $R$  for a particle which loses energy only through ionization and atomic excitation. Since  $dE/dx$  depends only on  $\beta$ ,  $R/M$  is a function of  $E/M$  or  $pc/M$ . In practice, range is a useful concept only for low-energy hadrons ( $R \lesssim \lambda_I$ , where  $\lambda_I$  is the nuclear interaction length), and for muons below a few hundred GeV (above which radiative effects dominate).  $R/M$  as a function of  $\beta\gamma = p/Mc$  is shown for a variety of materials in Fig. 33.4.

The mass scaling of  $dE/dx$  and range is valid for the electronic losses described by the Bethe equation, but not for radiative losses, relevant only for muons and pions.

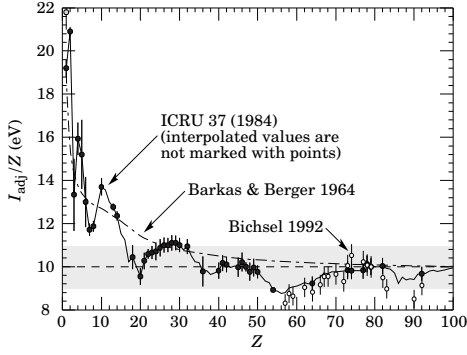
### 33.2.4. Mean excitation energy :

“The determination of the mean excitation energy is the principal non-trivial task in the evaluation of the Bethe stopping-power formula” [10]. Recommended values have varied substantially with time. Estimates based on experimental stopping-power measurements for protons, deuterons, and alpha particles and on oscillator-strength distributions and dielectric-response functions were given in ICRU 49 [4]. See also ICRU 37 [11]. These values, shown in Fig. 33.5, have since been widely used. Machine-readable versions can also be found [12].

### 33.2.5. Density effect :

As the particle energy increases, its electric field flattens and extends, so that the distant-collision contribution to Eq. (33.5) increases as  $\ln \beta\gamma$ . However, real media become polarized, limiting the field extension and effectively truncating this part of the logarithmic rise [2–8,15–16]. At very high energies,

$$\delta/2 \rightarrow \ln(\hbar\omega_p/I) + \ln \beta\gamma - 1/2, \quad (33.6)$$



**Figure 33.5:** Mean excitation energies (divided by  $Z$ ) as adopted by the ICRU [11]. Those based on experimental measurements are shown by symbols with error flags; the interpolated values are simply joined. The grey point is for liquid  $H_2$ ; the black point at 19.2 eV is for  $H_2$  gas. The open circles show more recent determinations by Bichsel [13]. The dash-dotted curve is from the approximate formula of Barkas [14] used in early editions of this Review.

where  $\delta(\beta\gamma)/2$  is the density effect correction introduced in Eq. (33.5) and  $\hbar\omega_p$  is the plasma energy defined in Table 33.1. A comparison with Eq. (33.5) shows that  $|dE/dx|$  then grows as  $\ln\beta\gamma$  rather than  $\ln\beta^2\gamma^2$ , and that the mean excitation energy  $I$  is replaced by the plasma energy  $\hbar\omega_p$ . The ionization stopping power as calculated with and without the density effect correction is shown in Fig. 33.1. Since the plasma frequency scales as the square root of the electron density, the correction is much larger for a liquid or solid than for a gas, as is illustrated by the examples in Fig. 33.2.

The density effect correction is usually computed using Sternheimer's parameterization [15]:

$$\delta(\beta\gamma) = \begin{cases} 2(\ln 10)x - \bar{C} & \text{if } x \geq x_1; \\ 2(\ln 10)x - \bar{C} + a(x_1 - x)^k & \text{if } x_0 \leq x < x_1; \\ 0 & \text{if } x < x_0 \text{ (nonconductors);} \\ \delta_0 10^{2(x-x_0)} & \text{if } x < x_0 \text{ (conductors)} \end{cases} \quad (33.7)$$

Here  $x = \log_{10} \eta = \log_{10}(p/Mc)$ .  $\bar{C}$  (the negative of the  $C$  used in Ref. 15) is obtained by equating the high-energy case of Eq. (33.7) with the limit given in Eq. (33.6). The other parameters are adjusted to give a best fit to the results of detailed calculations for momenta below  $Mc \exp(x_1)$ . Parameters for elements and nearly 200 compounds and mixtures of interest are published in a variety of places, notably in Ref. 16. A recipe for finding the coefficients for nontabulated materials is given by Sternheimer and Peierls [17], and is summarized in Ref. 5.

The remaining relativistic rise comes from the  $\beta^2\gamma$  growth of  $W_{\max}$ , which in turn is due to (rare) large energy transfers to a few electrons. When these events are excluded, the energy deposit in an absorbing layer approaches a constant value, the Fermi plateau (see Sec. 33.2.8 below). At even higher energies (*e.g.*,  $> 332$  GeV for muons in iron, and at a considerably higher energy for protons in iron), radiative effects are more important than ionization losses. These are especially relevant for high-energy muons, as discussed in Sec. 33.6.

### 33.2.6. Energy loss at low energies :

Shell corrections  $C/Z$  must be included in the square brackets of Eq. (33.5) [4,11,13,14] to correct for atomic binding having been neglected in calculating some of the contributions to Eq. (33.5). The Barkas form [14] was used in generating Fig. 33.1. For copper it contributes about 1% at  $\beta\gamma = 0.3$  (kinetic energy 6 MeV for a pion), and the correction decreases very rapidly with increasing energy.

Equation 33.2, and therefore Eq. (33.5), are based on a first-order Born approximation. Higher-order corrections, again important only at lower energies, are normally included by adding the “Bloch correction”  $z^2 L_2(\beta)$  inside the square brackets (Eq.(2.5) in [4]).

An additional “Barkas correction”  $z L_1(\beta)$  reduces the stopping power for a negative particle below that for a positive particle with the same mass and velocity. In a 1956 paper, Barkas *et al.* noted that

negative pions had a longer range than positive pions [6]. The effect has been measured for a number of negative/positive particle pairs, including a detailed study with antiprotons [18].

A detailed discussion of low-energy corrections to the Bethe formula is given in ICRU 49 [4]. When the corrections are properly included, the Bethe treatment is accurate to about 1% down to  $\beta \approx 0.05$ , or about 1 MeV for protons.

For  $0.01 < \beta < 0.05$ , there is no satisfactory theory. For protons, one usually relies on the phenomenological fitting formulae developed by Andersen and Ziegler [4,19]. As tabulated in ICRU 49 [4], the nuclear plus electronic proton stopping power in copper is  $113 \text{ MeV cm}^2 \text{ g}^{-1}$  at  $T = 10 \text{ keV}$  ( $\beta\gamma = 0.005$ ), rises to a maximum of  $210 \text{ MeV cm}^2 \text{ g}^{-1}$  at  $T \approx 120 \text{ keV}$  ( $\beta\gamma = 0.016$ ), then falls to  $118 \text{ MeV cm}^2 \text{ g}^{-1}$  at  $T = 1 \text{ MeV}$  ( $\beta\gamma = 0.046$ ). Above 0.5–1.0 MeV the corrected Bethe theory is adequate.

For particles moving more slowly than  $\approx 0.01c$  (more or less the velocity of the outer atomic electrons), Lindhard has been quite successful in describing electronic stopping power, which is proportional to  $\beta$  [20]. Finally, we note that at even lower energies, *e.g.*, for protons of less than several hundred eV, non-ionizing nuclear recoil energy loss dominates the total energy loss [4,20,21].

### 33.2.7. Energetic knock-on electrons ( $\delta$ rays) :

The distribution of secondary electrons with kinetic energies  $T \gg I$  is [2]

$$\frac{d^2 N}{dT dx} = \frac{1}{2} K z^2 \frac{Z}{A} \frac{1}{\beta^2} \frac{F(T)}{T^2} \quad (33.8)$$

for  $I \ll T \leq W_{\max}$ , where  $W_{\max}$  is given by Eq. (33.4). Here  $\beta$  is the velocity of the primary particle. The factor  $F$  is spin-dependent, but is about unity for  $T \ll W_{\max}$ . For spin-0 particles  $F(T) = (1 - \beta^2 T/W_{\max})$ ; forms for spins 1/2 and 1 are also given by Rossi [2] (Sec. 2.3, Eqns. 7 and 8). Additional formulae are given in Ref. 22. Equation (33.8) is inaccurate for  $T$  close to  $I$  [23].

$\delta$  rays of even modest energy are rare. For a  $\beta \approx 1$  particle, for example, on average only one collision with  $T_e > 10 \text{ keV}$  will occur along a path length of 90 cm of Ar gas [1].

A  $\delta$  ray with kinetic energy  $T_e$  and corresponding momentum  $p_e$  is produced at an angle  $\theta$  given by

$$\cos \theta = (T_e/p_e)(p_{\max}/W_{\max}), \quad (33.9)$$

where  $p_{\max}$  is the momentum of an electron with the maximum possible energy transfer  $W_{\max}$ .

### 33.2.8. Restricted energy loss rates for relativistic ionizing particles :

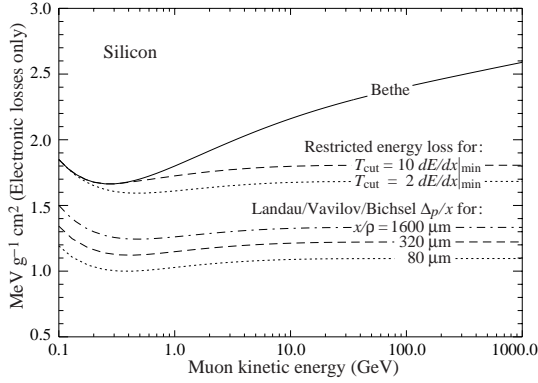
Further insight can be obtained by examining the mean energy deposit by an ionizing particle when energy transfers are restricted to  $T \leq W_{\text{cut}} \leq W_{\max}$ . The restricted energy loss rate is

$$\begin{aligned} -\frac{dE}{dx} \Big|_{T \leq W_{\text{cut}}} &= K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 W_{\text{cut}}}{I^2} \right. \\ &\quad \left. - \frac{\beta^2}{2} \left( 1 + \frac{W_{\text{cut}}}{W_{\max}} \right) - \frac{\delta}{2} \right]. \end{aligned} \quad (33.10)$$

This form approaches the normal Bethe function (Eq. (33.5)) as  $W_{\text{cut}} \rightarrow W_{\max}$ . It can be verified that the difference between Eq. (33.5) and Eq. (33.10) is equal to  $\int_{W_{\text{cut}}}^{W_{\max}} T (d^2 N/dT dx) dT$ , where  $d^2 N/dT dx$  is given by Eq. (33.8).

Since  $W_{\text{cut}}$  replaces  $W_{\max}$  in the argument of the logarithmic term of Eq. (33.5), the  $\beta\gamma$  term producing the relativistic rise in the close-collision part of  $dE/dx$  is replaced by a constant, and  $|dE/dx|_{T \leq W_{\text{cut}}}$  approaches the constant “Fermi plateau.” (The density effect correction  $\delta$  eliminates the explicit  $\beta\gamma$  dependence produced by the distant-collision contribution.) This behavior is illustrated in Fig. 33.6, where restricted loss rates for two examples of  $W_{\text{cut}}$  are shown in comparison with the full Bethe  $dE/dx$  and the Landau-Vavilov most probable energy loss (to be discussed in Sec. 33.2.9 below).

“Restricted energy loss” is cut at the total mean energy, not the single-collision energy above  $W_{\text{cut}}$ . It is of limited use. The most probable energy loss, discussed in the next Section, is far more useful in situations where single-particle energy loss is observed.



**Figure 33.6:** Bethe  $dE/dx$ , two examples of restricted energy loss, and the Landau most probable energy per unit thickness in silicon. The change of  $\Delta_p/x$  with thickness  $x$  illustrates its  $a \ln x + b$  dependence. Minimum ionization ( $dE/dx|_{\min}$ ) is  $1.664 \text{ MeV g}^{-1} \text{ cm}^2$ . Radiative losses are excluded. The incident particles are muons.

### 33.2.9. Fluctuations in energy loss :

For detectors of moderate thickness  $x$  (e.g. scintillators or LAr cells),\* the energy loss probability distribution  $f(\Delta; \beta\gamma, x)$  is adequately described by the highly-skewed Landau (or Landau-Vavilov) distribution [24,25]. The most probable energy loss is [26]

$$\Delta_p = \xi \left[ \ln \frac{2mc^2 \beta^2 \gamma^2}{I} + \ln \frac{\xi}{I} + j - \beta^2 - \delta(\beta\gamma) \right], \quad (33.11)$$

where  $\xi = (K/2) \langle Z/A \rangle z^2(x/\beta^2) \text{ MeV}$  for a detector with a thickness  $x$  in  $\text{g cm}^{-2}$ , and  $j = 0.200$  [26]. † While  $dE/dx$  is independent of thickness,  $\Delta_p/x$  scales as  $a \ln x + b$ . The density correction  $\delta(\beta\gamma)$  was not included in Landau's or Vavilov's work, but it was later included by Bichsel [26]. The high-energy behavior of  $\delta(\beta\gamma)$  (Eq. (33.6)) is such that

$$\Delta_p \xrightarrow{\beta\gamma \gg 100} \xi \left[ \ln \frac{2mc^2 \xi}{(\hbar\omega_p)^2} + j \right]. \quad (33.12)$$

Thus the Landau-Vavilov most probable energy loss, like the restricted energy loss, reaches a Fermi plateau. The Bethe  $dE/dx$  and Landau-Vavilov-Bichsel  $\Delta_p/x$  in silicon are shown as a function of muon energy in Fig. 33.6. The energy deposit in the  $1600 \mu\text{m}$  case is roughly the same as in a 3 mm thick plastic scintillator.

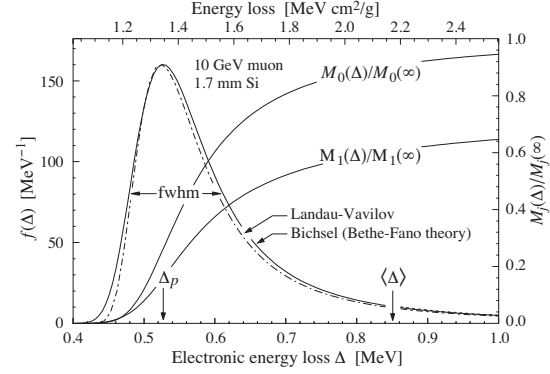
The distribution function for the energy deposit by a 10 GeV muon going through a detector of about this thickness is shown in Fig. 33.7. In this case the most probable energy loss is 62% of the mean ( $M_1(\langle\Delta\rangle)/M_1(\infty)$ ). Folding in experimental resolution displaces the peak of the distribution, usually toward a higher value. 90% of the collisions ( $M_1(\langle\Delta\rangle)/M_1(\infty)$ ) contribute to energy deposits below the mean. It is the very rare high-energy-transfer collisions, extending to  $W_{\max}$  at several GeV, that drives the mean into the tail of the distribution. The large weight of these rare events makes the mean of an experimental distribution consisting of a few hundred events subject to large fluctuations and sensitive to cuts. *The mean of the energy loss given by the Bethe equation, Eq. (33.5), is thus ill-defined experimentally and is not useful for describing energy loss by single particles.*‡ It rises as  $\ln \gamma$  because  $W_{\max}$  increases as  $\gamma$  at high energies.

\*  $G \lesssim 0.05\text{--}0.1$ , where  $G$  is given by Rossi [Ref. 2, Eq. 2.7(10)]. It is Vavilov's  $\kappa$  [25]. It is proportional to the absorber's thickness, and as such parameterizes the constants describing the Landau distribution. These are fairly insensitive to thickness for  $G \lesssim 0.1$ , the case for most detectors.

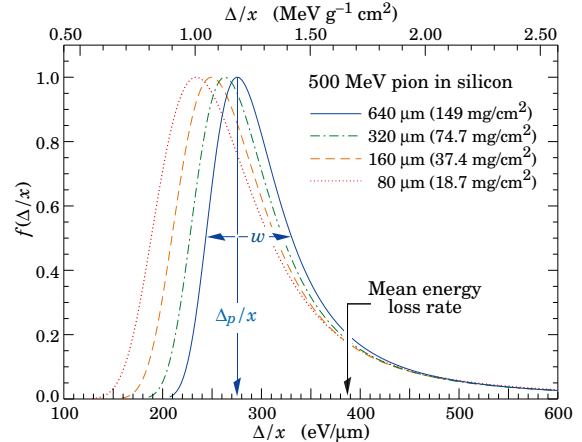
† Practical calculations can be expedited by using the tables of  $\delta$  and  $\beta$  from the text versions of the muon energy loss tables to be found at [pdg.lbl.gov/AtomicNuclearProperties](http://pdg.lbl.gov/AtomicNuclearProperties).

‡ Rossi [2], Talman [27], and others give somewhat different values for  $j$ . The most probable loss is not sensitive to its value.

§ It does find application in dosimetry, where only bulk deposit is relevant.



**Figure 33.7:** Electronic energy deposit distribution for a 10 GeV muon traversing 1.7 mm of silicon, the stopping power equivalent of about 0.3 cm of PVC scintillator [1,13,28]. The Landau-Vavilov function (dot-dashed) uses a Rutherford cross section without atomic binding corrections but with a kinetic energy transfer limit of  $W_{\max}$ . The solid curve was calculated using Bethe-Fano theory.  $M_0(\Delta)$  and  $M_1(\Delta)$  are the cumulative 0th moment (mean number of collisions) and 1st moment (mean energy loss) in crossing the silicon. (See Sec. 33.2.1. The fwhm of the Landau-Vavilov function is about  $4\xi$  for detectors of moderate thickness.  $\Delta_p$  is the most probable energy loss, and  $\langle\Delta\rangle$  divided by the thickness is the Bethe  $\langle dE/dx \rangle$ .)



**Figure 33.8:** Straggling functions in silicon for 500 MeV pions, normalized to unity at the most probable value  $\delta_p/x$ . The width  $w$  is the full width at half maximum.

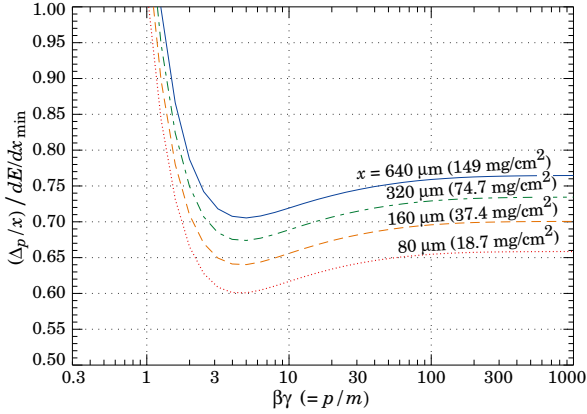
*The most probable energy loss should be used.*

A practical example: For muons traversing 0.25 inches of PVT plastic scintillator, the ratio of the most probable  $E$  loss rate to the mean loss rate via the Bethe equation is  $[0.69, 0.57, 0.49, 0.42, 0.38]$  for  $T_\mu = [0.01, 0.1, 1, 10, 100] \text{ GeV}$ . Radiative losses add less than 0.5% to the total mean energy deposit at 10 GeV, but add 7% at 100 GeV. The most probable  $E$  loss rate rises slightly beyond the minimum ionization energy, then is essentially constant.

The Landau distribution fails to describe energy loss in thin absorbers such as gas TPC cells [1] and Si detectors [26], as shown clearly in Fig. 1 of Ref. 1 for an argon-filled TPC cell. Also see Talman [27]. While  $\Delta_p/x$  may be calculated adequately with Eq. (33.11), the distributions are significantly wider than the Landau width  $w = 4\xi$  [Ref. 26, Fig. 15]. Examples for 500 MeV pions incident on thin silicon detectors are shown in Fig. 33.8. For very thick absorbers the distribution is less skewed but never approaches a Gaussian.

The most probable energy loss, scaled to the mean loss at minimum ionization, is shown in Fig. 33.9 for several silicon detector thicknesses.





**Figure 33.9:** Most probable energy loss in silicon, scaled to the mean loss of a minimum ionizing particle, 388 eV/μm (1.66 MeV g<sup>-1</sup>cm<sup>2</sup>).

### 33.2.10. Energy loss in mixtures and compounds :

A mixture or compound can be thought of as made up of thin layers of pure elements in the right proportion (Bragg additivity). In this case,

$$\left\langle \frac{dE}{dx} \right\rangle = \sum w_j \left\langle \frac{dE}{dx} \right\rangle_j, \quad (33.13)$$

where  $dE/dx|_j$  is the mean rate of energy loss (in MeV g cm<sup>-2</sup>) in the  $j$ th element. Eq. (33.5) can be inserted into Eq. (33.13) to find expressions for  $\langle Z/A \rangle$ ,  $\langle I \rangle$ , and  $\langle \delta \rangle$ ; for example,  $\langle Z/A \rangle = \sum w_j Z_j/A_j = \sum n_j Z_j / \sum n_j A_j$ . However,  $\langle I \rangle$  as defined this way is an underestimate, because in a compound electrons are more tightly bound than in the free elements, and  $\langle \delta \rangle$  as calculated this way has little relevance, because it is the electron density that matters. If possible, one uses the tables given in Refs. 16 and 29, that include effective excitation energies and interpolation coefficients for calculating the density effect correction for the chemical elements and nearly 200 mixtures and compounds. Otherwise, use the recipe for  $\delta$  given in Ref. 5 and 17, and calculate  $\langle I \rangle$  following the discussion in Ref. 10. (Note the “13%” rule!)

### 33.2.11. Ionization yields :

Physicists frequently relate total energy loss to the number of ion pairs produced near the particle’s track. This relation becomes complicated for relativistic particles due to the wandering of energetic knock-on electrons whose ranges exceed the dimensions of the fiducial volume. For a qualitative appraisal of the nonlocality of energy deposition in various media by such modestly energetic knock-on electrons, see Ref. 30. The mean local energy dissipation per local ion pair produced,  $W$ , while essentially constant for relativistic particles, increases at slow particle speeds [31]. For gases,  $W$  can be surprisingly sensitive to trace amounts of various contaminants [31]. Furthermore, ionization yields in practical cases may be greatly influenced by such factors as subsequent recombination [32].

## 33.3. Multiple scattering through small angles

A charged particle traversing a medium is deflected by many small-angle scatters. Most of this deflection is due to Coulomb scattering from nuclei as described by the Rutherford cross section. (However, for hadronic projectiles, the strong interactions also contribute to multiple scattering.) For many small-angle scatters the net scattering and displacement distributions are Gaussian via the central limit theorem. Less frequent “hard” scatters produce non-Gaussian tails. These Coulomb scattering distributions are well-represented by the theory of Molière [34]. Accessible discussions are given by Rossi [2] and Jackson [33], and exhaustive reviews have been published by Scott [35] and Motz *et al.* [36]. Experimental measurements have been published by Bichsel [37] (low energy protons) and by Shen *et al.* [38] (relativistic pions, kaons, and protons).\*

\* Shen *et al.*’s measurements show that Bethe’s simpler methods of including atomic electron effects agrees better with experiment than does Scott’s treatment.

If we define

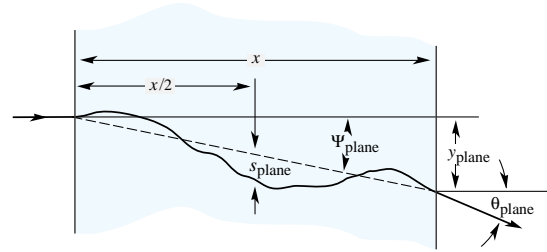
$$\theta_0 = \theta_{\text{plane}}^{\text{rms}} = \frac{1}{\sqrt{2}} \theta_{\text{space}}^{\text{rms}}, \quad (33.14)$$

then it is sufficient for many applications to use a Gaussian approximation for the central 98% of the projected angular distribution, with an rms width given by [39,40]

$$\theta_0 = \frac{13.6 \text{ MeV}}{\beta c p} z \sqrt{x/X_0} \left[ 1 + 0.038 \ln(x/X_0) \right]. \quad (33.15)$$

Here  $p$ ,  $\beta c$ , and  $z$  are the momentum, velocity, and charge number of the incident particle, and  $x/X_0$  is the thickness of the scattering medium in radiation lengths (defined below). This value of  $\theta_0$  is from a fit to Molière distribution for singly charged particles with  $\beta = 1$  for all  $Z$ , and is accurate to 11% or better for  $10^{-3} < x/X_0 < 100$ .

Eq. (33.15) describes scattering from a single material, while the usual problem involves the multiple scattering of a particle traversing many different layers and mixtures. Since it is from a fit to a Molière distribution, it is incorrect to add the individual  $\theta_0$  contributions in quadrature; the result is systematically too small. It is much more accurate to apply Eq. (33.15) once, after finding  $x$  and  $X_0$  for the combined scatterer.



**Figure 33.10:** Quantities used to describe multiple Coulomb scattering. The particle is incident in the plane of the figure.

The nonprojected (space) and projected (plane) angular distributions are given approximately by [34]

$$\frac{1}{2\pi \theta_0^2} \exp \left( -\frac{\theta_{\text{space}}^2}{2\theta_0^2} \right) d\Omega, \quad (33.16)$$

$$\frac{1}{\sqrt{2\pi} \theta_0} \exp \left( -\frac{\theta_{\text{plane}}^2}{2\theta_0^2} \right) d\theta_{\text{plane}}, \quad (33.17)$$

where  $\theta$  is the deflection angle. In this approximation,  $\theta_{\text{space}}^2 \approx (\theta_{\text{plane},x}^2 + \theta_{\text{plane},y}^2)$ , where the  $x$  and  $y$  axes are orthogonal to the direction of motion, and  $d\Omega \approx d\theta_{\text{plane},x} d\theta_{\text{plane},y}$ . Deflections into  $\theta_{\text{plane},x}$  and  $\theta_{\text{plane},y}$  are independent and identically distributed.

Fig. 33.10 shows these and other quantities sometimes used to describe multiple Coulomb scattering. They are

$$\psi_{\text{plane}}^{\text{rms}} = \frac{1}{\sqrt{3}} \theta_{\text{plane}}^{\text{rms}} = \frac{1}{\sqrt{3}} \theta_0, \quad (33.18)$$

$$y_{\text{plane}}^{\text{rms}} = \frac{1}{\sqrt{3}} x \theta_{\text{plane}}^{\text{rms}} = \frac{1}{\sqrt{3}} x \theta_0, \quad (33.19)$$

$$s_{\text{plane}}^{\text{rms}} = \frac{1}{4\sqrt{3}} x \theta_{\text{plane}}^{\text{rms}} = \frac{1}{4\sqrt{3}} x \theta_0. \quad (33.20)$$

All the quantitative estimates in this section apply only in the limit of small  $\theta_{\text{plane}}^{\text{rms}}$  and in the absence of large-angle scatters. The random variables  $s$ ,  $\psi$ ,  $y$ , and  $\theta$  in a given plane are correlated. Obviously,  $y \approx x\psi$ . In addition,  $y$  and  $\theta$  have the correlation coefficient  $\rho_{y\theta} = \sqrt{3}/2 \approx 0.87$ . For Monte Carlo generation of a joint  $(y_{\text{plane}}, \theta_{\text{plane}})$  distribution, or for other calculations, it may be most convenient to work with independent Gaussian random variables ( $z_1, z_2$ ) with mean zero and variance one, and then set

$$y_{\text{plane}} = z_1 x \theta_0 (1 - \rho_{y\theta}^2)^{1/2} / \sqrt{3} + z_2 \rho_{y\theta} x \theta_0 / \sqrt{3} \quad (33.21)$$

$$= z_1 x \theta_0 / \sqrt{12} + z_2 x \theta_0 / 2; \quad (33.22)$$

$$\theta_{\text{plane}} = z_2 \theta_0. \quad (33.23)$$

Note that the second term for  $y_{\text{plane}}$  equals  $x \theta_{\text{plane}}/2$  and represents the displacement that would have occurred had the deflection  $\theta_{\text{plane}}$  all occurred at the single point  $x/2$ .

For heavy ions the multiple Coulomb scattering has been measured and compared with various theoretical distributions [41].

### 33.4. Photon and electron interactions in matter

At low energies electrons and positrons primarily lose energy by ionization, although other processes (Møller scattering, Bhabha scattering,  $e^+$  annihilation) contribute, as shown in Fig. 33.11. While ionization loss rates rise logarithmically with energy, bremsstrahlung losses rise nearly linearly (fractional loss is nearly independent of energy), and dominates above the critical energy (Sec. 33.4.4 below), a few tens of MeV in most materials

#### 33.4.1. Collision energy losses by $e^\pm$ :

Stopping power differs somewhat for electrons and positrons, and both differ from stopping power for heavy particles because of the kinematics, spin, charge, and the identity of the incident electron with the electrons that it ionizes. Complete discussions and tables can be found in Refs. 10, 11, and 29.

For electrons, large energy transfers to atomic electrons (taken as free) are described by the Møller cross section. From Eq. (33.4), the maximum energy transfer in a single collision should be the entire kinetic energy,  $W_{\max} = m_e c^2(\gamma - 1)$ , but because the particles are identical, the maximum is half this,  $W_{\max}/2$ . (The results are the same if the transferred energy is  $\epsilon$  or if the transferred energy is  $W_{\max} - \epsilon$ . The stopping power is by convention calculated for the faster of the two emerging electrons.) The first moment of the Møller cross section [22] (divided by  $dx$ ) is the stopping power:

$$\left\langle -\frac{dE}{dx} \right\rangle = \frac{1}{2} K \frac{Z}{A} \frac{1}{\beta^2} \left[ \ln \frac{m_e c^2 \beta^2 \gamma^2 \{m_e c^2 (\gamma - 1)/2\}}{I^2} + (1 - \beta^2) - \frac{2\gamma - 1}{\gamma^2} \ln 2 + \frac{1}{8} \left( \frac{\gamma - 1}{\gamma} \right)^2 - \delta \right] \quad (33.24)$$

The logarithmic term can be compared with the logarithmic term in the Bethe equation (Eq. (33.2)) by substituting  $W_{\max} = m_e c^2(\gamma - 1)/2$ . The two forms differ by  $\ln 2$ .

Electron-positron scattering is described by the fairly complicated Bhabha cross section [22]. There is no identical particle problem, so  $W_{\max} = m_e c^2(\gamma - 1)$ . The first moment of the Bhabha equation yields

$$\left\langle -\frac{dE}{dx} \right\rangle = \frac{1}{2} K \frac{Z}{A} \frac{1}{\beta^2} \left[ \ln \frac{m_e c^2 \beta^2 \gamma^2 \{m_e c^2 (\gamma - 1)\}}{2I^2} + 2 \ln 2 - \frac{\beta^2}{12} \left( 23 + \frac{14}{\gamma + 1} + \frac{10}{(\gamma + 1)^2} + \frac{4}{(\gamma + 1)^3} \right) - \delta \right]. \quad (33.25)$$

Following ICRU 37 [11], the density effect correction  $\delta$  has been added to Uehling's equations [22] in both cases.

For heavy particles, shell corrections were developed assuming that the projectile is equivalent to a perturbing potential whose center moves with constant velocity. This assumption has no sound theoretical basis for electrons. The authors of ICRU 37 [11] estimated the possible error in omitting it by assuming the correction was twice as great as for a proton of the same velocity. At  $T = 10$  keV, the error was estimated to be  $\approx 2\%$  for water,  $\approx 9\%$  for Cu, and  $\approx 21\%$  for Au.

As shown in Fig. 33.11, stopping powers for  $e^-$ ,  $e^+$ , and heavy particles are not dramatically different. In silicon, the minimum value for electrons is  $1.50 \text{ MeV cm}^2/\text{g}$  (at  $\gamma = 3.3$ ); for positrons,  $1.46 \text{ MeV cm}^2/\text{g}$  (at  $\gamma = 3.7$ ), and for muons,  $1.66 \text{ MeV cm}^2/\text{g}$  (at  $\gamma = 3.58$ ).

#### 33.4.2. Radiation length:

High-energy electrons predominantly lose energy in matter by bremsstrahlung, and high-energy photons by  $e^+e^-$  pair production. The characteristic amount of matter traversed for these related interactions is called the radiation length  $X_0$ , usually measured in  $\text{g cm}^{-2}$ . It is both (a) the mean distance over which a high-energy electron loses all but  $1/e$  of its energy by bremsstrahlung, and (b)  $\frac{7}{9}$  of the mean free path for pair production by a high-energy photon [42]. It is also the appropriate scale length for describing high-energy electromagnetic cascades.  $X_0$  has been calculated and tabulated by Y.S. Tsai [43]:

$$\frac{1}{X_0} = 4\alpha r_e^2 \frac{N_A}{A} \left\{ Z^2 [L_{\text{rad}} - f(Z)] + Z L'_{\text{rad}} \right\}. \quad (33.26)$$

For  $A = 1 \text{ g mol}^{-1}$ ,  $4\alpha r_e^2 N_A/A = (716.408 \text{ g cm}^{-2})^{-1}$ .  $L_{\text{rad}}$  and  $L'_{\text{rad}}$  are given in Table 33.2. The function  $f(Z)$  is an infinite sum, but for elements up to uranium can be represented to 4-place accuracy by

$$f(Z) = a^2 \left[ (1 + a^2)^{-1} + 0.20206 - 0.0369 a^2 + 0.0083 a^4 - 0.002 a^6 \right], \quad (33.27)$$

where  $a = \alpha Z$  [44].

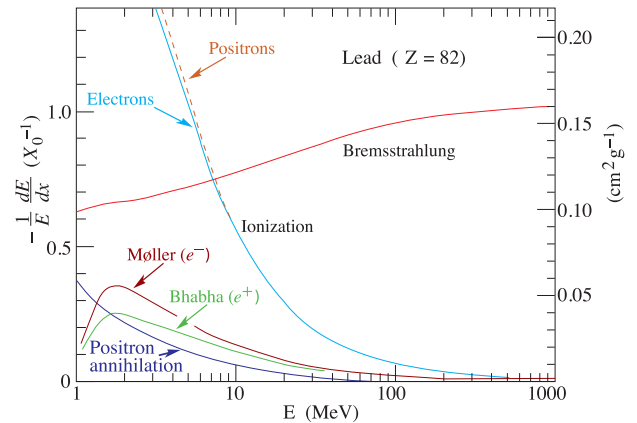
**Table 33.2:** Tsai's  $L_{\text{rad}}$  and  $L'_{\text{rad}}$  for use in calculating the radiation length in an element using Eq. (33.26).

Element	$Z$	$L_{\text{rad}}$	$L'_{\text{rad}}$
H	1	5.31	6.144
He	2	4.79	5.621
Li	3	4.74	5.805
Be	4	4.71	5.924
Others	$> 4$	$\ln(184.15 Z^{-1/3})$	$\ln(1194 Z^{-2/3})$

The radiation length in a mixture or compound may be approximated by

$$1/X_0 = \sum w_j/X_j, \quad (33.28)$$

where  $w_j$  and  $X_j$  are the fraction by weight and the radiation length for the  $j$ th element.



**Figure 33.11:** Fractional energy loss per radiation length in lead as a function of electron or positron energy. Electron (positron) scattering is considered as ionization when the energy loss per collision is below  $0.255 \text{ MeV}$ , and as Møller (Bhabha) scattering when it is above. Adapted from Fig. 3.2 from Messel and Crawford, *Electron-Photon Shower Distribution Function Tables for Lead, Copper, and Air Absorbers*, Pergamon Press, 1970. Messel and Crawford use  $X_0(\text{Pb}) = 5.82 \text{ g/cm}^2$ , but we have modified the figures to reflect the value given in the Table of Atomic and Nuclear Properties of Materials ( $X_0(\text{Pb}) = 6.37 \text{ g/cm}^2$ ).

#### 33.4.3. Bremsstrahlung energy loss by $e^\pm$ :

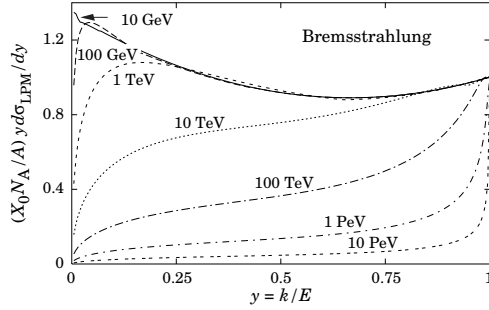
At very high energies and except at the high-energy tip of the bremsstrahlung spectrum, the cross section can be approximated in the “complete screening case” as [43]

$$\frac{d\sigma}{dk} = (1/k) 4\alpha r_e^2 \left\{ \left( \frac{4}{3} - \frac{4}{3}y + y^2 \right) [Z^2 (L_{\text{rad}} - f(Z)) + Z L'_{\text{rad}}] + \frac{1}{9} (1 - y) (Z^2 + Z) \right\}, \quad (33.29)$$

where  $y = k/E$  is the fraction of the electron's energy transferred to the radiated photon. At small  $y$  (the “infrared limit”) the term on the second line ranges from 1.7% (low  $Z$ ) to 2.5% (high  $Z$ ) of the total. If it is ignored and the first line simplified with the definition of  $X_0$  given in Eq. (33.26), we have

$$\frac{d\sigma}{dk} = \frac{A}{X_0 N_A k} \left( \frac{4}{3} - \frac{4}{3}y + y^2 \right). \quad (33.30)$$

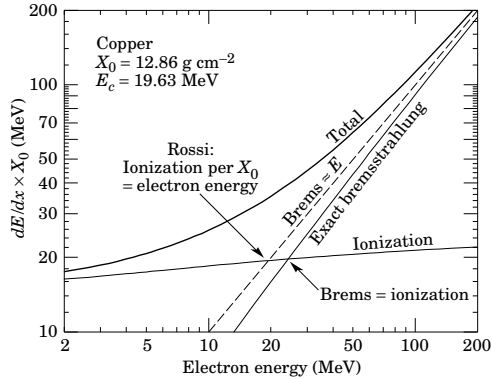
This cross section (times  $k$ ) is shown by the top curve in Fig. 33.12.



**Figure 33.12:** The normalized bremsstrahlung cross section  $k d\sigma_{LPM}/dk$  in lead versus the fractional photon energy  $y = k/E$ . The vertical axis has units of photons per radiation length.

This formula is accurate except in near  $y = 1$ , where screening may become incomplete, and near  $y = 0$ , where the infrared divergence is removed by the interference of bremsstrahlung amplitudes from nearby scattering centers (the LPM effect) [45,46] and dielectric suppression [47,48]. These and other suppression effects in bulk media are discussed in Sec. 33.4.6.

With decreasing energy ( $E \lesssim 10$  GeV) the high- $y$  cross section drops and the curves become rounded as  $y \rightarrow 1$ . Curves of this familiar shape can be seen in Rossi [2] (Figs. 2.11.2,3); see also the review by Koch & Motz [49].



**Figure 33.13:** Two definitions of the critical energy  $E_c$ .

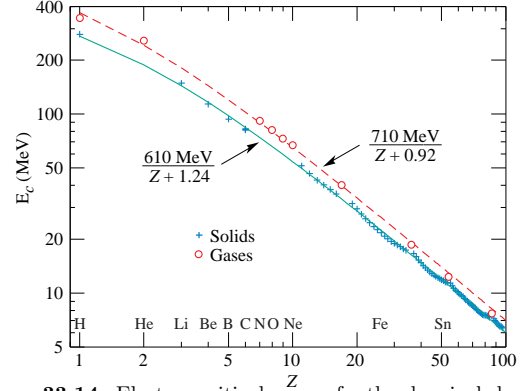
Except at these extremes, and still in the complete-screening approximation, the number of photons with energies between  $k_{\min}$  and  $k_{\max}$  emitted by an electron travelling a distance  $d \ll X_0$  is

$$N_\gamma = \frac{d}{X_0} \left[ \frac{4}{3} \ln \left( \frac{k_{\max}}{k_{\min}} \right) - \frac{4(k_{\max} - k_{\min})}{3E} + \frac{k_{\max}^2 - k_{\min}^2}{2E^2} \right]. \quad (33.31)$$

#### 33.4.4. Critical energy :

An electron loses energy by bremsstrahlung at a rate nearly proportional to its energy, while the ionization loss rate varies only logarithmically with the electron energy. The *critical energy*  $E_c$  is sometimes defined as the energy at which the two loss rates are equal [50]. Among alternate definitions is that of Rossi [2], who defines the critical energy as the energy at which the ionization loss per radiation length is equal to the electron energy. Equivalently, it is the same as the first definition with the approximation  $|dE/dx|_{\text{brems}} \approx E/X_0$ . This form has been found to describe transverse electromagnetic shower development more accurately (see below). These definitions are illustrated in the case of copper in Fig. 33.13.

The accuracy of approximate forms for  $E_c$  has been limited by the failure to distinguish between gases and solid or liquids, where there is a substantial difference in ionization at the relevant energy because of the density effect. We distinguish these two cases in Fig. 33.14.



**Figure 33.14:** Electron critical energy for the chemical elements, using Rossi's definition [2]. The fits shown are for solids and liquids (solid line) and gases (dashed line). The rms deviation is 2.2% for the solids and 4.0% for the gases. (Computed with code supplied by A. Fassó.)

Fits were also made with functions of the form  $a/(Z+b)^\alpha$ , but  $\alpha$  was found to be essentially unity. Since  $E_c$  also depends on  $A$ ,  $I$ , and other factors, such forms are at best approximate.

Values of  $E_c$  for both electrons and positrons in more than 300 materials can be found at [pdg.lbl.gov/AtomicNuclearProperties](http://pdg.lbl.gov/AtomicNuclearProperties).

#### 33.4.5. Energy loss by photons :

Contributions to the photon cross section in a light element (carbon) and a heavy element (lead) are shown in Fig. 33.15. At low energies it is seen that the photoelectric effect dominates, although Compton scattering, Rayleigh scattering, and photonuclear absorption also contribute. The photoelectric cross section is characterized by discontinuities (absorption edges) as thresholds for photoionization of various atomic levels are reached. Photon attenuation lengths for a variety of elements are shown in Fig. 33.18, and data for  $30 \text{ eV} < k < 100 \text{ GeV}$  for all elements are available from the web pages given in the caption. Here  $k$  is the photon energy.

The increasing domination of pair production as the energy increases is shown in Fig. 33.16. Using approximations similar to those used to obtain Eq. (33.30), Tsai's formula for the differential cross section [43] reduces to

$$\frac{d\sigma}{dx} = \frac{A}{X_0 N_A} \left[ 1 - \frac{4}{3} x(1-x) \right] \quad (33.32)$$

in the complete-screening limit valid at high energies. Here  $x = E/k$  is the fractional energy transfer to the pair-produced electron (or positron), and  $k$  is the incident photon energy. The cross section is very closely related to that for bremsstrahlung, since the Feynman diagrams are variants of one another. The cross section is of necessity symmetric between  $x$  and  $1-x$ , as can be seen by the solid curve in Fig. 33.17. See the review by Motz, Olsen, & Koch for a more detailed treatment [53].

Eq. (33.32) may be integrated to find the high-energy limit for the total  $e^+e^-$  pair-production cross section:

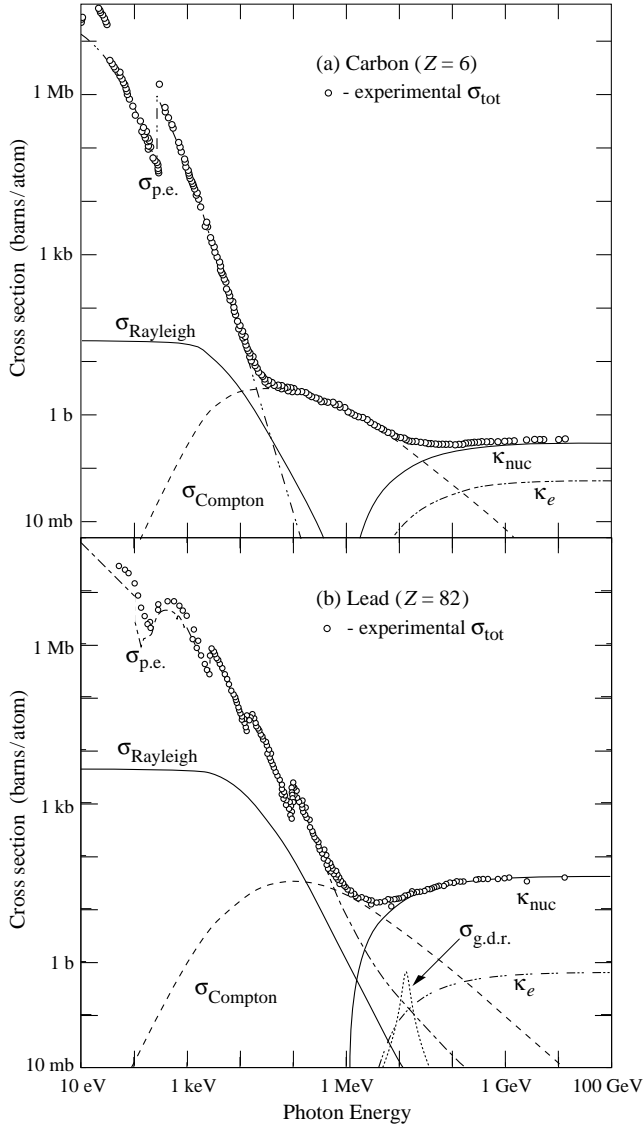
$$\sigma = \frac{7}{9} (A/X_0 N_A). \quad (33.33)$$

Equation (33.33) is accurate to within a few percent down to energies as low as 1 GeV, particularly for high- $Z$  materials.

#### 33.4.6. Bremsstrahlung and pair production at very high energies :

At ultrahigh energies, Eqns. 33.29–33.33 will fail because of quantum mechanical interference between amplitudes from different scattering centers. Since the longitudinal momentum transfer to a given center is small ( $\propto k/E(E-k)$ , in the case of bremsstrahlung), the interaction is spread over a comparatively long distance called the formation length ( $\propto E(E-k)/k$ ) via the uncertainty principle. In alternate language, the formation length is the distance over which the highly relativistic electron and the photon “split apart.” The interference is usually destructive. Calculations of the “Landau-Pomeranchuk-Migdal” (LPM) effect may be made semi-classically based on the average multiple scattering, or more rigorously using a quantum transport approach [45,46].





**Figure 33.15:** Photon total cross sections as a function of energy in carbon and lead, showing the contributions of different processes [51]:

- $\sigma_{p.e.}$  = Atomic photoelectric effect (electron ejection, photon absorption)
- $\sigma_{\text{Rayleigh}}$  = Rayleigh (coherent) scattering—atom neither ionized nor excited
- $\sigma_{\text{Compton}}$  = Incoherent scattering (Compton scattering off an electron)
- $\kappa_{\text{nuc}}$  = Pair production, nuclear field
- $\kappa_e$  = Pair production, electron field
- $\sigma_{g.d.r.}$  = Photonuclear interactions, most notably the Giant Dipole Resonance [52]. In these interactions, the target nucleus is broken up.

Original figures through the courtesy of John H. Hubbell (NIST).

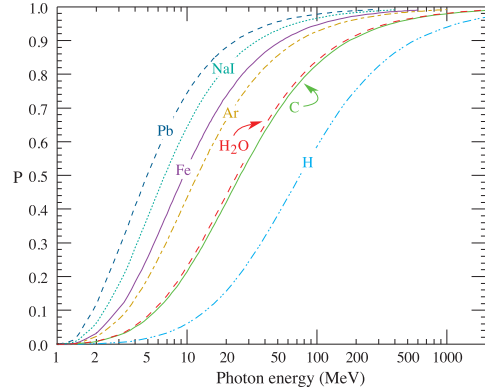
In amorphous media, bremsstrahlung is suppressed if the photon energy  $k$  is less than  $E^2/(E + E_{LPM})$  [46], where\*

$$E_{LPM} = \frac{(m_e c^2)^2 \alpha X_0}{4\pi \hbar c \rho} = (7.7 \text{ TeV/cm}) \times \frac{X_0}{\rho}. \quad (33.34)$$

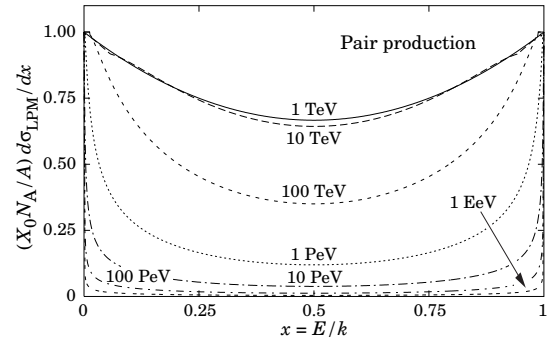
Since physical distances are involved,  $X_0/\rho$ , in cm, appears. The energy-weighted bremsstrahlung spectrum for lead,  $k d\sigma_{LPM}/dk$ ,

\* This definition differs from that of Ref. 54 by a factor of two.  $E_{LPM}$  scales as the 4th power of the mass of the incident particle, so that  $E_{LPM} = (1.4 \times 10^{10} \text{ TeV/cm}) \times X_0/\rho$  for a muon.

is shown in Fig. 33.12. With appropriate scaling by  $X_0/\rho$ , other materials behave similarly.

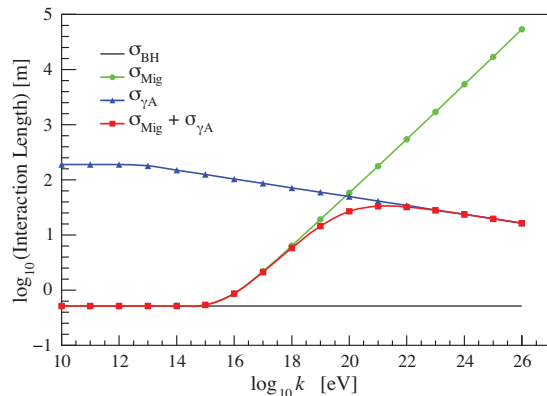


**Figure 33.16:** Probability  $P$  that a photon interaction will result in conversion to an  $e^+e^-$  pair. Except for a few-percent contribution from photonuclear absorption around 10 or 20 MeV, essentially all other interactions in this energy range result in Compton scattering off an atomic electron. For a photon attenuation length  $\lambda$  (Fig. 33.18), the probability that a given photon will produce an electron pair (without first Compton scattering) in thickness  $t$  of absorber is  $P[1 - \exp(-t/\lambda)]$ .

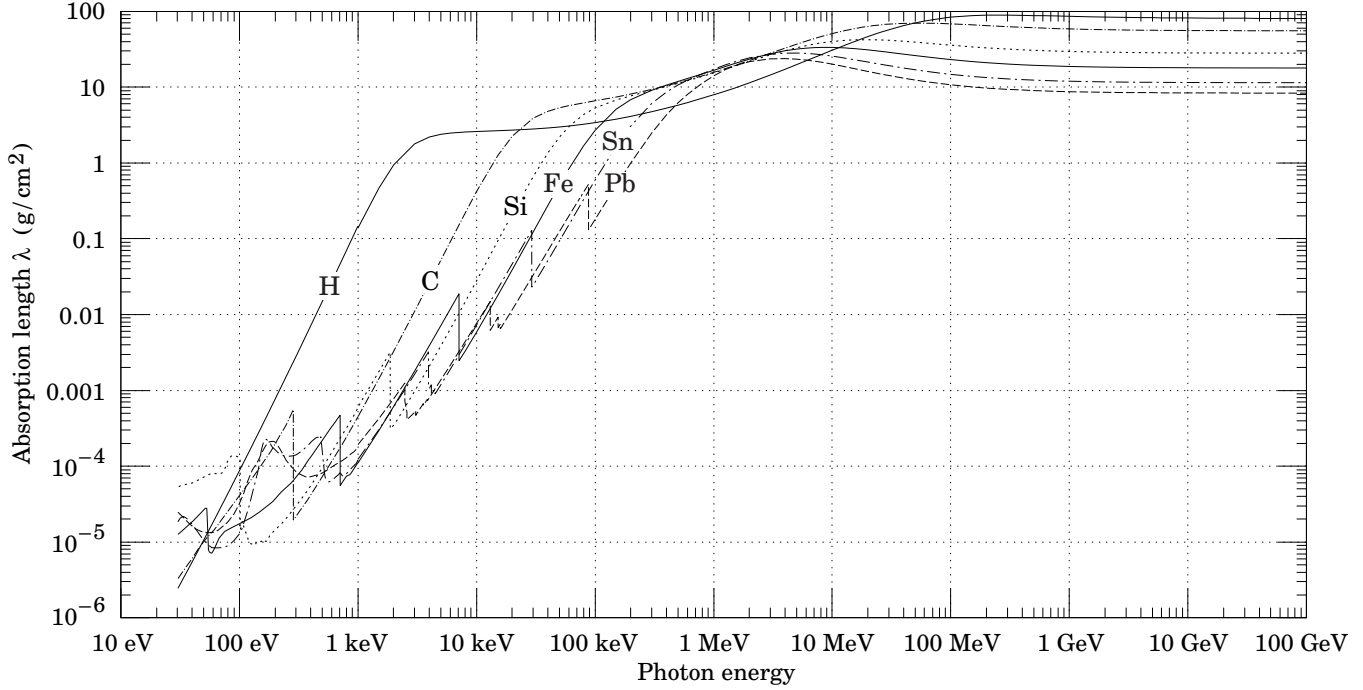


**Figure 33.17:** The normalized pair production cross section  $d\sigma_{LPM}/dy$ , versus fractional electron energy  $x = E/k$ .

For photons, pair production is reduced for  $E(k - E) > k E_{LPM}$ . The pair-production cross sections for different photon energies are shown in Fig. 33.17.

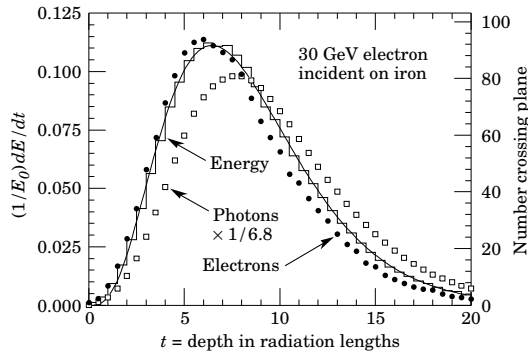


**Figure 33.19:** Interaction length for a photon in ice as a function of photon energy for the Bethe-Heitler (BH), LPM (Mig) and photonuclear ( $\gamma A$ ) cross sections [56]. The Bethe-Heitler interaction length is  $9X_0/7$ , and  $X_0$  is 0.393 m in ice.



**Figure 33.18:** The photon mass attenuation length (or mean free path)  $\lambda = 1/(\mu/\rho)$  for various elemental absorbers as a function of photon energy. The mass attenuation coefficient is  $\mu/\rho$ , where  $\rho$  is the density. The intensity  $I$  remaining after traversal of thickness  $t$  (in mass/unit area) is given by  $I = I_0 \exp(-t/\lambda)$ . The accuracy is a few percent. For a chemical compound or mixture,  $1/\lambda_{\text{eff}} \approx \sum_{\text{elements}} w_Z/\lambda_Z$ , where  $w_Z$  is the proportion by weight of the element with atomic number  $Z$ . The processes responsible for attenuation are given in Fig. 33.11. Since coherent processes are included, not all these processes result in energy deposition. The data for  $30 \text{ eV} < E < 1 \text{ keV}$  are obtained from [http://www-cxro.lbl.gov/optical\\_constants](http://www-cxro.lbl.gov/optical_constants) (courtesy of Eric M. Gullikson, LBNL). The data for  $1 \text{ keV} < E < 100 \text{ GeV}$  are from <http://physics.nist.gov/PhysRefData>, through the courtesy of John H. Hubbell (NIST).

If  $k \ll E$ , several additional mechanisms can also produce suppression. When the formation length is long, even weak factors can perturb the interaction. For example, the emitted photon can coherently forward scatter off of the electrons in the media. Because of this, for  $k < \omega_p E/m_e \sim 10^{-4}$ , bremsstrahlung is suppressed by a factor  $(k m_e / \omega_p E)^2$  [48]. Magnetic fields can also suppress bremsstrahlung.



**Figure 33.20:** An EGS4 simulation of a 30 GeV electron-induced cascade in iron. The histogram shows fractional energy deposition per radiation length, and the curve is a gamma-function fit to the distribution. Circles indicate the number of electrons with total energy greater than 1.5 MeV crossing planes at  $X_0/2$  intervals (scale on right) and the squares the number of photons with  $E \geq 1.5 \text{ MeV}$  crossing the planes (scaled down to have same area as the electron distribution).

In crystalline media, the situation is more complicated, with coherent enhancement or suppression possible. The cross section depends on the electron and photon energies and the angles between the particle direction and the crystalline axes [55].

### 33.4.7. Photonuclear and electronuclear interactions at still higher energies :

At still higher photon and electron energies, where the bremsstrahlung and pair production cross-sections are heavily suppressed by the LPM effect, photonuclear and electronuclear interactions predominate over electromagnetic interactions.

At photon energies above about  $10^{20} \text{ eV}$ , for example, photons usually interact hadronically. The exact cross-over energy depends on the model used for the photonuclear interactions. These processes are illustrated in Fig. 33.19. At still higher energies ( $\gtrsim 10^{23} \text{ eV}$ ), photonuclear interactions can become coherent, with the photon interaction spread over multiple nuclei. Essentially, the photon coherently converts to a  $\rho^0$ , in a process that is somewhat similar to kaon regeneration [56].

Similar processes occur for electrons. As electron energies increase and the LPM effect suppresses bremsstrahlung, electronuclear interactions become more important. At energies above  $10^{21} \text{ eV}$ , these electronuclear interactions dominate electron energy loss [56].

### 33.5. Electromagnetic cascades

When a high-energy electron or photon is incident on a thick absorber, it initiates an electromagnetic cascade as pair production and bremsstrahlung generate more electrons and photons with lower energy. The longitudinal development is governed by the high-energy part of the cascade, and therefore scales as the radiation length in the material. Electron energies eventually fall below the critical energy, and then dissipate their energy by ionization and excitation rather than by the generation of more shower particles. In describing shower behavior, it is therefore convenient to introduce the scale variables

$$t = x/X_0, \quad y = E/E_c, \quad (33.35)$$

so that distance is measured in units of radiation length and energy in units of critical energy. Longitudinal profiles from an EGS4 [57] simulation of a 30 GeV electron-induced cascade in iron are shown in Fig. 33.20. The number of particles crossing a plane (very close to Rossi's  $\Pi$  function [2]) is sensitive to the cutoff energy, here chosen as a total energy of 1.5 MeV for both electrons and photons. The electron

number falls off more quickly than energy deposition. This is because, with increasing depth, a larger fraction of the cascade energy is carried by photons. Exactly what a calorimeter measures depends on the device, but it is not likely to be exactly any of the profiles shown. In gas counters it may be very close to the electron number, but in glass Cherenkov detectors and other devices with “thick” sensitive regions it is closer to the energy deposition (total track length). In such detectors the signal is proportional to the “detectable” track length  $T_d$ , which is in general less than the total track length  $T$ . Practical devices are sensitive to electrons with energy above some detection threshold  $E_d$ , and  $T_d = T F(E_d/E_c)$ . An analytic form for  $F(E_d/E_c)$  obtained by Rossi [2] is given by Fabjan in Ref. 58; see also Amaldi [59].

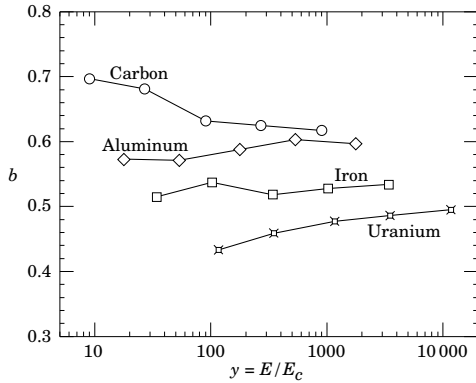
The mean longitudinal profile of the energy deposition in an electromagnetic cascade is reasonably well described by a gamma distribution [60]:

$$\frac{dE}{dt} = E_0 b \frac{(bt)^{a-1} e^{-bt}}{\Gamma(a)} \quad (33.36)$$

The maximum  $t_{\max}$  occurs at  $(a-1)/b$ . We have made fits to shower profiles in elements ranging from carbon to uranium, at energies from 1 GeV to 100 GeV. The energy deposition profiles are well described by Eq. (33.36) with

$$t_{\max} = (a-1)/b = 1.0 \times (\ln y + C_j), \quad j = e, \gamma, \quad (33.37)$$

where  $C_e = -0.5$  for electron-induced cascades and  $C_\gamma = +0.5$  for photon-induced cascades. To use Eq. (33.36), one finds  $(a-1)/b$  from Eq. (33.37) and Eq. (33.35), then finds  $a$  either by assuming  $b \approx 0.5$  or by finding a more accurate value from Fig. 33.21. The results are very similar for the electron number profiles, but there is some dependence on the atomic number of the medium. A similar form for the electron number maximum was obtained by Rossi in the context of his “Approximation B,” [2] (see Fabjan’s review in Ref. 58), but with  $C_e = -1.0$  and  $C_\gamma = -0.5$ ; we regard this as superseded by the EGS4 result.



**Figure 33.21:** Fitted values of the scale factor  $b$  for energy deposition profiles obtained with EGS4 for a variety of elements for incident electrons with  $1 \leq E_0 \leq 100$  GeV. Values obtained for incident photons are essentially the same.

The “shower length”  $X_s = X_0/b$  is less conveniently parameterized, since  $b$  depends upon both  $Z$  and incident energy, as shown in Fig. 33.21. As a corollary of this  $Z$  dependence, the number of electrons crossing a plane near shower maximum is underestimated using Rossi’s approximation for carbon and seriously overestimated for uranium. Essentially the same  $b$  values are obtained for incident electrons and photons. For many purposes it is sufficient to take  $b \approx 0.5$ .

The length of showers initiated by ultra-high energy photons and electrons is somewhat greater than at lower energies since the first or first few interaction lengths are increased via the mechanisms discussed above.

The gamma function distribution is very flat near the origin, while the EGS4 cascade (or a real cascade) increases more rapidly. As a result Eq. (33.36) fails badly for about the first two radiation lengths; it was necessary to exclude this region in making fits.

Because fluctuations are important, Eq. (33.36) should be used only in applications where average behavior is adequate. Grindhammer *et al.* have developed fast simulation algorithms in which the variance and correlation of  $a$  and  $b$  are obtained by fitting Eq. (33.36) to individually simulated cascades, then generating profiles for cascades using  $a$  and  $b$  chosen from the correlated distributions [61].

The transverse development of electromagnetic showers in different materials scales fairly accurately with the *Molière radius*  $R_M$ , given by [62,63]

$$R_M = X_0 E_s / E_c, \quad (33.38)$$

where  $E_s \approx 21$  MeV (Table 33.1), and the Rossi definition of  $E_c$  is used.

In a material containing a weight fraction  $w_j$  of the element with critical energy  $E_{cj}$  and radiation length  $X_j$ , the Molière radius is given by

$$\frac{1}{R_M} = \frac{1}{E_s} \sum \frac{w_j E_{cj}}{X_j}. \quad (33.39)$$

Measurements of the lateral distribution in electromagnetic cascades are shown in Refs. 62 and 63. On the average, only 10% of the energy lies outside the cylinder with radius  $R_M$ . About 99% is contained inside of  $3.5R_M$ , but at this radius and beyond composition effects become important and the scaling with  $R_M$  fails. The distributions are characterized by a narrow core, and broaden as the shower develops. They are often represented as the sum of two Gaussians, and Grindhammer [61] describes them with the function

$$f(r) = \frac{2r R^2}{(r^2 + R^2)^2}, \quad (33.40)$$

where  $R$  is a phenomenological function of  $x/X_0$  and  $\ln E$ .

At high enough energies, the LPM effect (Sec. 33.4.6) reduces the cross sections for bremsstrahlung and pair production, and hence can cause significant elongation of electromagnetic cascades [46].

### 33.6. Muon energy loss at high energy

At sufficiently high energies, radiative processes become more important than ionization for all charged particles. For muons and pions in materials such as iron, this “critical energy” occurs at several hundred GeV. (There is no simple scaling with particle mass, but for protons the “critical energy” is much, much higher.) Radiative effects dominate the energy loss of energetic muons found in cosmic rays or produced at the newest accelerators. These processes are characterized by small cross sections, hard spectra, large energy fluctuations, and the associated generation of electromagnetic and (in the case of photonuclear interactions) hadronic showers [64–72]. As a consequence, at these energies the treatment of energy loss as a uniform and continuous process is for many purposes inadequate.

It is convenient to write the average rate of muon energy loss as [73]

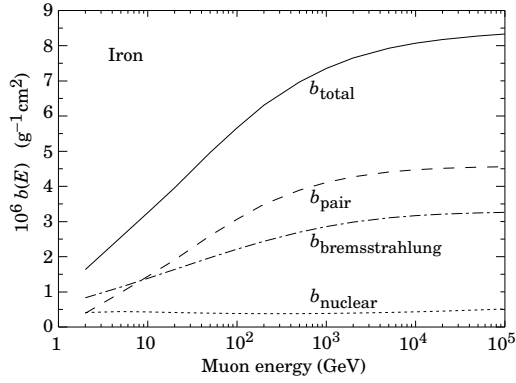
$$-dE/dx = a(E) + b(E)E. \quad (33.41)$$

Here  $a(E)$  is the ionization energy loss given by Eq. (33.5), and  $b(E)$  is the sum of  $e^+e^-$  pair production, bremsstrahlung, and photonuclear contributions. To the approximation that these slowly-varying functions are constant, the mean range  $x_0$  of a muon with initial energy  $E_0$  is given by

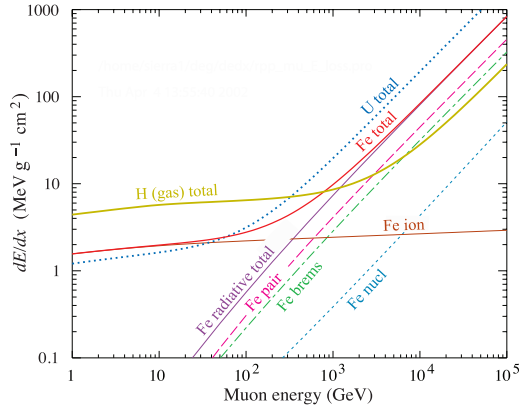
$$x_0 \approx (1/b) \ln(1 + E_0/E_{\mu c}), \quad (33.42)$$

where  $E_{\mu c} = a/b$ . Fig. 33.22 shows contributions to  $b(E)$  for iron. Since  $a(E) \approx 0.002$  GeV  $g^{-1} cm^2$ ,  $b(E)E$  dominates the energy loss above several hundred GeV, where  $b(E)$  is nearly constant. The rates of energy loss for muons in hydrogen, uranium, and iron are shown in Fig. 33.23 [5].

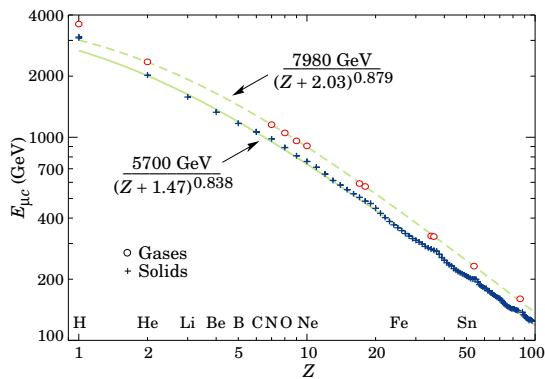
The “muon critical energy”  $E_{\mu c}$  can be defined more exactly as the energy at which radiative and ionization losses are equal, and can be found by solving  $E_{\mu c} = a(E_{\mu c})/b(E_{\mu c})$ . This definition corresponds to the solid-line intersection in Fig. 33.13, and is different from the Rossi definition we used for electrons. It serves the same function: below  $E_{\mu c}$  ionization losses dominate, and above  $E_{\mu c}$  radiative effects dominate. The dependence of  $E_{\mu c}$  on atomic number  $Z$  is shown in Fig. 33.24.



**Figure 33.22:** Contributions to the fractional energy loss by muons in iron due to  $e^+e^-$  pair production, bremsstrahlung, and photonuclear interactions, as obtained from Groom *et al.* [5] except for post-Born corrections to the cross section for direct pair production from atomic electrons.



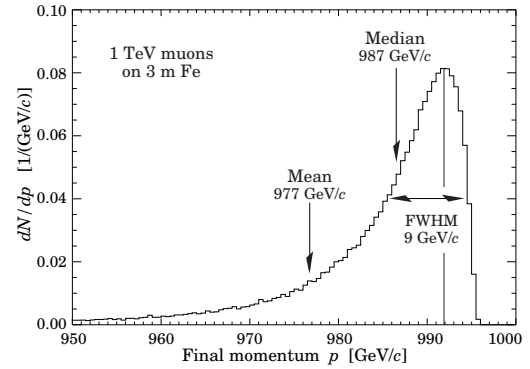
**Figure 33.23:** The average energy loss of a muon in hydrogen, iron, and uranium as a function of muon energy. Contributions to  $dE/dx$  in iron from ionization and the processes shown in Fig. 33.22 are also shown.



**Figure 33.24:** Muon critical energy for the chemical elements, defined as the energy at which radiative and ionization energy loss rates are equal [5]. The equality comes at a higher energy for gases than for solids or liquids with the same atomic number because of a smaller density effect reduction of the ionization losses. The fits shown in the figure exclude hydrogen. Alkali metals fall 3–4% above the fitted function, while most other solids are within 2% of the function. Among the gases the worst fit is for radon (2.7% high).

The radiative cross sections are expressed as functions of the fractional energy loss  $\nu$ . The bremsstrahlung cross section goes roughly as  $1/\nu$  over most of the range, while for the pair production case the distribution goes as  $\nu^{-3}$  to  $\nu^{-2}$  [74]. “Hard” losses are

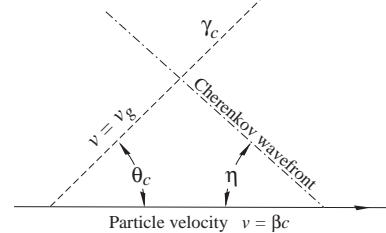
therefore more probable in bremsstrahlung, and in fact energy losses due to pair production may very nearly be treated as continuous. The simulated [72] momentum distribution of an incident 1 TeV/c muon beam after it crosses 3 m of iron is shown in Fig. 33.25. The most probable loss is 8 GeV, or  $3.4 \text{ MeV g}^{-1}\text{cm}^2$ . The full width at half maximum is 9 GeV/c, or 0.9%. The radiative tail is almost entirely due to bremsstrahlung, although most of the events in which more than 10% of the incident energy lost experienced relatively hard photonuclear interactions. The latter can exceed detector resolution [75], necessitating the reconstruction of lost energy. Tables in Ref. 5 list the stopping power as  $9.82 \text{ MeV g}^{-1}\text{cm}^2$  for a 1 TeV muon, so that the mean loss should be 23 GeV ( $\approx 23 \text{ GeV/c}$ ), for a final momentum of 977 GeV/c, far below the peak. This agrees with the indicated mean calculated from the simulation. Electromagnetic and hadronic cascades in detector materials can obscure muon tracks in detector planes and reduce tracking efficiency [76].



**Figure 33.25:** The momentum distribution of 1 TeV/c muons after traversing 3 m of iron as calculated with the MARS15 Monte Carlo code [72] by S.I. Striganov [5].

### 33.7. Cherenkov and transition radiation [33,77,78]

A charged particle radiates if its velocity is greater than the local phase velocity of light (Cherenkov radiation) or if it crosses suddenly from one medium to another with different optical properties (transition radiation). Neither process is important for energy loss, but both are used in high-energy and cosmic-ray physics detectors.



**Figure 33.26:** Cherenkov light emission and wavefront angles. In a dispersive medium,  $\theta_c + \eta \neq 90^\circ$ .

#### 33.7.1. Optical Cherenkov radiation :

The angle  $\theta_c$  of Cherenkov radiation, relative to the particle's direction, for a particle with velocity  $\beta c$  in a medium with index of refraction  $n$  is

$$\begin{aligned} \cos \theta_c &= (1/n\beta) \\ \text{or } \tan \theta_c &= \sqrt{\beta^2 n^2 - 1} \\ &\approx \sqrt{2(1 - 1/n\beta)} \quad \text{for small } \theta_c, \text{ e.g. in gases.} \end{aligned} \quad (33.43)$$

The threshold velocity  $\beta_t$  is  $1/n$ , and  $\gamma_t = 1/(1 - \beta_t^2)^{1/2}$ . Therefore,  $\beta_t \gamma_t = 1/(2\delta + \delta^2)^{1/2}$ , where  $\delta = n - 1$ . Values of  $\delta$  for various commonly used gases are given as a function of pressure and wavelength in Ref. 79. For values at atmospheric pressure, see Table 6.1. Data for other commonly used materials are given in Ref. 80.

Practical Cherenkov radiator materials are dispersive. Let  $\omega$  be the photon's frequency, and let  $k = 2\pi/\lambda$  be its wavenumber. The photons propagate at the group velocity  $v_g = d\omega/dk = c/[n(\omega) + \omega(dn/d\omega)]$ . In a non-dispersive medium, this simplifies to  $v_g = c/n$ .

In his classical paper, Tamm [81] showed that for dispersive media the radiation is concentrated in a thin conical shell whose vertex is at the moving charge, and whose opening half-angle  $\eta$  is given by

$$\cot \eta = \left[ \frac{d}{d\omega} (\omega \tan \theta_c) \right]_{\omega_0} = \left[ \tan \theta_c + \beta^2 \omega n(\omega) \frac{dn}{d\omega} \cot \theta_c \right]_{\omega_0}, \quad (33.44)$$

where  $\omega_0$  is the central value of the small frequency range under consideration. (See Fig. 33.26.) This cone has a opening half-angle  $\eta$ , and, unless the medium is non-dispersive ( $dn/d\omega = 0$ ),  $\theta_c + \eta \neq 90^\circ$ . The Cherenkov wavefront 'sideslips' along with the particle [82]. This effect has timing implications for ring imaging Cherenkov counters [83], but it is probably unimportant for most applications.

The number of photons produced per unit path length of a particle with charge  $ze$  and per unit energy interval of the photons is

$$\frac{d^2 N}{dE dx} = \frac{\alpha z^2}{\hbar c} \sin^2 \theta_c = \frac{\alpha^2 z^2}{r_e m_e c^2} \left( 1 - \frac{1}{\beta^2 n^2(E)} \right) \approx 370 \sin^2 \theta_c(E) \text{ eV}^{-1} \text{ cm}^{-1} \quad (z = 1), \quad (33.45)$$

or, equivalently,

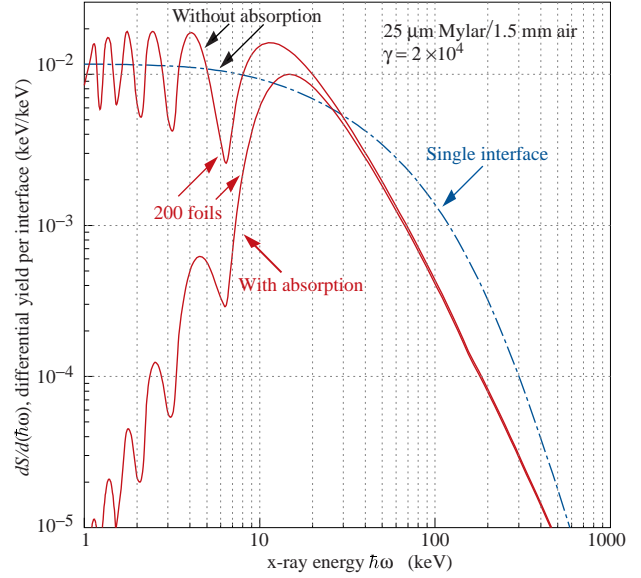
$$\frac{d^2 N}{dx d\lambda} = \frac{2\pi \alpha z^2}{\lambda^2} \left( 1 - \frac{1}{\beta^2 n^2(\lambda)} \right). \quad (33.46)$$

The index of refraction  $n$  is a function of photon energy  $E = \hbar\omega$ , as is the sensitivity of the transducer used to detect the light. For practical use, Eq. (33.45) must be multiplied by the transducer response function and integrated over the region for which  $\beta n(\omega) > 1$ . Further details are given in the discussion of Cherenkov detectors in the Particle Detectors section (Sec. 34 of this Review).

When two particles are close together (lateral separation  $\lesssim 1$  wavelength), the electromagnetic fields from the particles may add coherently, affecting the Cherenkov radiation. Because of their opposite charges, the radiation from an  $e^+e^-$  pair at close separation is suppressed compared to two independent leptons [84].

### 33.7.2. Coherent radio Cherenkov radiation :

Coherent Cherenkov radiation is produced by many charged particles with a non-zero net charge moving through matter on an approximately common "wavefront"—for example, the electrons and positrons in a high-energy electromagnetic cascade. The signals can be visible above backgrounds for shower energies as low as  $10^{17}$  eV; see Sec. 35.3.3 for more details. The phenomenon is called the Askaryan effect [85]. Near the end of a shower, when typical particle energies are below  $E_c$  (but still relativistic), a charge imbalance develops. Photons can Compton-scatter atomic electrons, and positrons can annihilate with atomic electrons to contribute even more photons which can in turn Compton scatter. These processes result in a roughly 20% excess of electrons over positrons in a shower. The net negative charge leads to coherent radio Cherenkov emission. The radiation includes a component from the decelerating charges (as in bremsstrahlung). Because the emission is coherent, the electric field strength is proportional to the shower energy, and the signal power increases as its square. The electric field strength also increases linearly with frequency, up to a maximum frequency determined by the lateral spread of the shower. This cutoff occurs at about 1 GHz in ice, and scales inversely with the Moliere radius. At low frequencies, the radiation is roughly isotropic, but, as the frequency rises toward the cutoff frequency, the radiation becomes increasingly peaked around the Cherenkov angle. The radiation is linearly polarized in the plane containing the shower axis and the photon direction. A measurement of the signal polarization can be used to help determine the shower direction. The characteristics of this radiation have been



**Figure 33.27:** X-ray photon energy spectra for a radiator consisting of 200 25  $\mu\text{m}$  thick foils of Mylar with 1.5 mm spacing in air (solid lines) and for a single surface (dashed line). Curves are shown with and without absorption. Adapted from Ref. 88.

nically demonstrated in a series of experiments at SLAC [86]. A detailed discussion of the radiation can be found in Ref. 87.

### 33.7.3. Transition radiation :

The energy radiated when a particle with charge  $ze$  crosses the boundary between vacuum and a medium with plasma frequency  $\omega_p$  is

$$I = \alpha z^2 \gamma \hbar \omega_p / 3, \quad (33.47)$$

where

$$\hbar \omega_p = \sqrt{4\pi N_e r_e^3 m_e c^2 / \alpha} = \sqrt{\rho \text{ (in g/cm}^3\text{)} \langle Z/A \rangle \times 28.81 \text{ eV}}. \quad (33.48)$$

For styrene and similar materials,  $\hbar \omega_p \approx 20$  eV; for air it is 0.7 eV.

The number spectrum  $dN_\gamma/d(\hbar\omega)$  diverges logarithmically at low energies and decreases rapidly for  $\hbar\omega/\gamma\hbar\omega_p > 1$ . About half the energy is emitted in the range  $0.1 \leq \hbar\omega/\gamma\hbar\omega_p \leq 1$ . Inevitable absorption in a practical detector removes the divergence. For a particle with  $\gamma = 10^3$ , the radiated photons are in the soft x-ray range 2 to 40 keV. The  $\gamma$  dependence of the emitted energy thus comes from the hardening of the spectrum rather than from an increased quantum yield.

The number of photons with energy  $\hbar\omega > \hbar\omega_0$  is given by the answer to problem 13.15 in Ref. 33,

$$N_\gamma(\hbar\omega > \hbar\omega_0) = \frac{\alpha z^2}{\pi} \left[ \left( \ln \frac{\gamma \hbar \omega_p}{\hbar \omega_0} - 1 \right)^2 + \frac{\pi^2}{12} \right], \quad (33.49)$$

within corrections of order  $(\hbar\omega_0/\gamma\hbar\omega_p)^2$ . The number of photons above a fixed energy  $\hbar\omega_0 \ll \gamma\hbar\omega_p$  thus grows as  $(\ln \gamma)^2$ , but the number above a fixed fraction of  $\gamma\hbar\omega_p$  (as in the example above) is constant. For example, for  $\hbar\omega > \gamma\hbar\omega_p/10$ ,  $N_\gamma = 2.519 \alpha z^2 / \pi = 0.59\% \times z^2$ .

The particle stays "in phase" with the x ray over a distance called the formation length,  $d(\omega) = (2c/\omega)(1/\gamma^2 + \theta^2 + \omega_p^2/\omega^2)^{-1}$ . Most of the radiation is produced in this distance. Here  $\theta$  is the x-ray emission angle, characteristically  $1/\gamma$ . For  $\theta = 1/\gamma$  the formation length has a maximum at  $d(\gamma\omega_p/\sqrt{2}) = \gamma c/\sqrt{2}\omega_p$ . In practical situations it is tens of  $\mu\text{m}$ .

Since the useful x-ray yield from a single interface is low, in practical detectors it is enhanced by using a stack of  $N$  foil radiators—foils  $L$  thick, where  $L$  is typically several formation lengths—separated by gas-filled gaps. The amplitudes at successive interfaces interfere to cause oscillations about the single-interface spectrum. At increasing frequencies above the position of the last interference maximum ( $L/d(\omega) = \pi/2$ ), the formation zones, which have opposite phase, overlap more and more and the spectrum saturates,  $dI/d\omega$  approaching zero as  $L/d(\omega) \rightarrow 0$ . This is illustrated in Fig. 33.27 for a realistic detector configuration.

For regular spacing of the layers fairly complicated analytic solutions for the intensity have been obtained [88,89]. Although one might expect the intensity of coherent radiation from the stack of foils to be proportional to  $N^2$ , the angular dependence of the formation length conspires to make the intensity  $\propto N$ .

#### References:

- H. Bichsel, Nucl. Instrum. Methods **A562**, 154 (2006).
- B. Rossi, *High Energy Particles*, Prentice-Hall, Inc., Englewood Cliffs, NJ, 1952.
- H.A. Bethe, *Zur Theorie des Durchgangs schneller Korpuskularstrahlen durch Materie*, H. Bethe, Ann. Phys. **5**, 325 (1930).
- "Stopping Powers and Ranges for Protons and Alpha Particles," ICRU Report No. 49 (1993); tables and graphs of these data are available at <http://physics.nist.gov/PhysRefData/>.
- D.E. Groom, N.V. Mokhov, and S.I. Striganov, "Muon stopping-power and range tables: 10 MeV–100 TeV," Atomic Data and Nuclear Data Tables **78**, 183–356 (2001). Since submission of this paper it has become likely that post-Born corrections to the direct pair production cross section should be made. Code used to make Figs. 33.22–33.24 included these corrections [D.Yu. Ivanov *et al.*, Phys. Lett. **B442**, 453 (1998)]. The effect is negligible except at high  $Z$ . (It is less than 1% for iron.); More extensive printable and machine-readable tables are given at <http://pdg.lbl.gov/AtomicNuclearProperties/>.
- W.H. Barkas, W. Birnbaum, and F.M. Smith, Phys. Rev. **101**, 778 (1956).
- J. Lindhard and A. H. Sørensen, Phys. Rev. **A53**, 2443 (1996).
- U. Fano, Ann. Rev. Nucl. Sci. **13**, 1 (1963).
- J.D. Jackson, Phys. Rev. **D59**, 017301 (1999).
- S.M. Seltzer and M.J. Berger, Int. J. of Applied Rad. **33**, 1189 (1982).
- "Stopping Powers for Electrons and Positrons," ICRU Report No. 37 (1984).
- <http://physics.nist.gov/PhysRefData/XrayMassCoef/tab1.html>.
- H. Bichsel, Phys. Rev. **A46**, 5761 (1992).
- W.H. Barkas and M.J. Berger, *Tables of Energy Losses and Ranges of Heavy Charged Particles*, NASA-SP-3013 (1964).
- R.M. Sternheimer, Phys. Rev. **88**, 851 (1952).
- R.M. Sternheimer, S.M. Seltzer, and M.J. Berger, "The Density Effect for the Ionization Loss of Charged Particles in Various Substances," Atomic Data and Nuclear Data Tables **30**, 261 (1984). Minor errors are corrected in Ref. 5. Chemical composition for the tabulated materials is given in Ref. 10.
- R.M. Sternheimer and R.F. Peierls, Phys. Rev. **B3**, 3681 (1971).
- S.P. Møller *et al.*, Phys. Rev. **A56**, 2930 (1997).
- H.H. Andersen and J.F. Ziegler, *Hydrogen: Stopping Powers and Ranges in All Elements*. Vol. 3 of *The Stopping and Ranges of Ions in Matter* (Pergamon Press 1977).
- J. Lindhard, Kgl. Danske Videnskab. Selskab, Mat.-Fys. Medd. **28**, No. 8 (1954); J. Lindhard, M. Scharff, and H.E. Schiøtt, Kgl. Danske Videnskab. Selskab, Mat.-Fys. Medd. **33**, No. 14 (1963).
- J.F. Ziegler, J.F. Biersac, and U. Littmark, *The Stopping and Range of Ions in Solids*, Pergamon Press 1985.
- E.A. Uehling, Ann. Rev. Nucl. Sci. **4**, 315 (1954) (For heavy particles with unit charge, but  $e^\pm$  cross sections and stopping powers are also given).
- N.F. Mott and H.S.W. Massey, *The Theory of Atomic Collisions*, Oxford Press, London, 1965.
- L.D. Landau, J. Exp. Phys. (USSR) **8**, 201 (1944).
- P.V. Vavilov, Sov. Phys. JETP **5**, 749 (1957).
- H. Bichsel, Rev. Mod. Phys. **60**, 663 (1988).
- R. Talman, Nucl. Instrum. Methods **159**, 189 (1979).
- H. Bichsel, Ch. 87 in the Atomic, Molecular and Optical Physics Handbook, G.W.F. Drake, editor (Am. Inst. Phys. Press, Woodbury NY, 1996).
- S.M. Seltzer and M.J. Berger, Int. J. of Applied Rad. **35**, 665 (1984). This paper corrects and extends the results of Ref. 10.
- L.V. Spencer "Energy Dissipation by Fast Electrons," Nat'l Bureau of Standards Monograph No. 1 (1959).
- "Average Energy Required to Produce an Ion Pair," ICRU Report No. 31 (1979).
- N. Hadley *et al.*, "List of Poisoning Times for Materials," Lawrence Berkeley Lab Report TPC-LBL-79-8 (1981).
- J.D. Jackson, *Classical Electrodynamics*, 3rd edition, (John Wiley and Sons, New York, 1998).
- H.A. Bethe, Phys. Rev. **89**, 1256 (1953).
- W.T. Scott, Rev. Mod. Phys. **35**, 231 (1963).
- J.W. Motz, H. Olsen, and H.W. Koch, Rev. Mod. Phys. **36**, 881 (1964).
- H. Bichsel, Phys. Rev. **112**, 182 (1958).
- G. Shen *et al.*, (Phys. Rev. **D20**, 1584 (1979)).
- V.L. Highland, Nucl. Instrum. Methods **129**, 497 (1975); Nucl. Instrum. Methods **161**, 171 (1979).
- G.R. Lynch and O.I. Dahl, Nucl. Instrum. Methods **B58**, 6 (1991). Eq. (33.15) is Eq. 12 from this paper.
- M. Wong *et al.*, Med. Phys. **17**, 163 (1990).
- E. Segrè, *Nuclei and Particles*, New York, Benjamin (1964) p. 65 ff.
- Y.S. Tsai, Rev. Mod. Phys. **46**, 815 (1974).
- H. Davies, H.A. Bethe, and L.C. Maximon, Phys. Rev. **93**, 788 (1954).
- L.D. Landau and I.J. Pomeranchuk, Dokl. Akad. Nauk. SSSR **92**, 535 (1953); **92**, 735 (1953). These papers are available in English in L. Landau, *The Collected Papers of L.D. Landau*, Pergamon Press, 1965; A.B. Migdal, Phys. Rev. **103**, 1811 (1956).
- S. Klein, Rev. Mod. Phys. **71**, 1501 (1999).
- M.L. Ter-Mikaelian, SSSR **94**, 1033 (1954); M.L. Ter-Mikaelian, *High Energy Electromagnetic Processes in Condensed Media* (John Wiley and Sons, New York, 1972).
- P. Anthony *et al.*, Phys. Rev. Lett. **76**, 3550 (1996).
- H.W. Koch and J.W. Motz, Rev. Mod. Phys. **31**, 920 (1959).
- M.J. Berger and S.M. Seltzer, "Tables of Energy Losses and Ranges of Electrons and Positrons," National Aeronautics and Space Administration Report NASA-SP-3012 (Washington DC 1964).
- Data from J.H. Hubbell, H. Gimm, and I. Øverbø, J. Phys. Chem. Ref. Data **9**, 1023 (1980); parameters for  $\sigma_{g.d.r.}$  from A. Veyssiere *et al.*, Nucl. Phys. **A159**, 561 (1970). Curves for these and other elements, compounds, and mixtures may be obtained from <http://physics.nist.gov/PhysRefData>. The photon total cross section is approximately flat for at least two decades beyond the energy range shown.
- B.L. Berman and S.C. Fultz, Rev. Mod. Phys. **47**, 713 (1975).
- J.W. Motz, H.A. Olsen, and H.W. Koch, Rev. Mod. Phys. **41**, 581 (1969).
- P. Anthony *et al.*, Phys. Rev. Lett. **75**, 1949 (1995).
- U.I. Uggerhøj, Rev. Mod. Phys. **77**, 1131 (2005).
- L. Gerhardt and S.R. Klein, Phys. Rev. **D82**, 074017 (2010).
- W.R. Nelson, H. Hirayama, and D.W.O. Rogers, "The EGS4 Code System," SLAC-265, Stanford Linear Accelerator Center (Dec. 1985).
- Experimental Techniques in High Energy Physics*, ed. T. Ferbel (Addison-Wesley, Menlo Park CA 1987).
- U. Amaldi, Phys. Scripta **23**, 409 (1981).
- E. Longo and I. Sestili, Nucl. Instrum. Methods **128**, 283 (1975).
- G. Grindhammer *et al.*, in *Proceedings of the Workshop on Calorimetry for the Supercollider*, Tuscaloosa, AL, March 13–17, 1989, edited by R. Donaldson and M.G.D. Gilchriese (World Scientific, Teaneck, NJ, 1989), p. 151.
- W.R. Nelson *et al.*, Phys. Rev. **149**, 201 (1966).
- G. Bathow *et al.*, Nucl. Phys. **B20**, 592 (1970).



64. H.A. Bethe and W. Heitler, Proc. Royal Soc. London **A146**, 83 (1934);  
H.A. Bethe, Proc. Cambridge Phil. Soc. **30**, 542 (1934).
65. A.A. Petrukhin and V.V. Shestakov, Can. J. Phys. **46**, S377 (1968).
66. V.M. Galitskii and S.R. Kel'ner, Sov. Phys. JETP **25**, 948 (1967).
67. S.R. Kel'ner and Yu.D. Kotov, Sov. J. Nucl. Phys. **7**, 237 (1968).
68. R.P. Kokoulin and A.A. Petrukhin, in *Proceedings of the International Conference on Cosmic Rays*, Hobart, Australia, August 16–25, 1971, Vol. **4**, p. 2436.
69. A.I. Nikishov, Sov. J. Nucl. Phys. **27**, 677 (1978).
70. Y.M. Andreev *et al.*, Phys. Atom. Nucl. **57**, 2066 (1994).
71. L.B. Bezrukov and E.V. Bugaev, Sov. J. Nucl. Phys. **33**, 635 (1981).
72. N.V. Mokhov, "The MARS Code System User's Guide," Fermilab-FN-628 (1995);  
N.V. Mokhov *et al.*, Radiation Protection and Dosimetry, vol. 116, part 2, pp. 99 (2005);  
Fermilab-Conf-04/053 (2004);  
N.V. Mokhov *et al.*, in *Proc. of Intl. Conf. on Nuclear Data for Science and Tech*, (Santa Fe, NM, 2004), AIP Conf. Proc. 769, part 2, p. 1618;  
Fermilab-Conf-04/269-AD (2004);  
<http://www-ap.fnal.gov/MARS/>.
73. P.H. Barrett *et al.*, Rev. Mod. Phys. **24**, 133 (1952).
74. A. Van Ginneken, Nucl. Instrum. Methods **A251**, 21 (1986).
75. U. Becker *et al.*, Nucl. Instrum. Methods **A253**, 15 (1986).
76. J.J. Eastman and S.C. Loken, in *Proceedings of the Workshop on Experiments, Detectors, and Experimental Areas for the Supercollider*, Berkeley, CA, July 7–17, 1987, edited by R. Donaldson and M.G.D. Gilchriese (World Scientific, Singapore, 1988), p. 542.
77. *Methods of Experimental Physics*, L.C.L. Yuan and C.-S. Wu, editors, Academic Press, 1961, Vol. 5A, p. 163.
78. W.W.M. Allison and P.R.S. Wright, "The Physics of Charged Particle Identification:  $dE/dx$ , Cherenkov Radiation, and Transition Radiation," p. 371 in *Experimental Techniques in High Energy Physics*, T. Ferbel, editor, (Addison-Wesley 1987).
79. E.R. Hayes, R.A. Schluter, and A. Tamosaitis, "Index and Dispersion of Some Cherenkov Counter Gases," ANL-6916 (1964).
80. T. Ypsilantis, "Particle Identification at Hadron Colliders," CERN-EP/89-150 (1989), or ECFA 89-124, **2** 661 (1989).
81. I. Tamm, J. Phys. U.S.S.R., **1**, 439 (1939).
82. H. Motz and L.I. Schiff, Am. J. Phys. **21**, 258 (1953).
83. B.N. Ratcliff, Nucl. Instrum. Methods **A502**, 211 (2003).
84. S.K. Mandal, S.R. Klein, and J. D. Jackson, Phys. Rev. **D72**, 093003 (2005).
85. G.A. Askaryan, Sov. Phys. JETP **14**, 441 (1962).
86. P.W. Gorham *et al.*, Phys. Rev. **D72**, 023002 (2005).
87. E. Zas, F. Halzen, and T. Stanev, Phys. Rev. D **45**, 362 (1992).
88. M.L. Cherry, Phys. Rev. **D10**, 3594 (1974);  
M.L. Cherry, Phys. Rev. **D17**, 2245 (1978).
89. B. Dolgoshein, Nucl. Instrum. Methods **A326**, 434 (1993).

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## 34.1. Introduction

This review summarizes the detector technologies employed at accelerator particle physics experiments. Several of these detectors are also used in a non-accelerator context and examples of such applications will be provided. The detector techniques which are specific to non-accelerator particle physics experiments are the subject of Chap. 35. More detailed discussions of detectors and their underlying physics can be found in books by Ferbel [1], Kleinknecht [2], Knoll [3], Green [4], Leroy & Rancoita [5], and Grupen [6].

In Table 34.1 are given typical resolutions and deadtimes of common charged particle detectors. The quoted numbers are usually based on typical devices, and should be regarded only as rough approximations for new designs. The spatial resolution refers to the intrinsic detector resolution, i.e. without multiple scattering. We note that analog detector readout can provide better spatial resolution than digital readout by measuring the deposited charge in neighboring channels. Quoted ranges attempt to be representative of both possibilities. The time resolution is defined by how accurately the time at which a particle crossed the detector can be determined. The deadtime is the minimum separation in time between two resolved hits on the same channel. Typical performance of calorimetry and particle identification are provided in the relevant sections below.

**Table 34.1:** Typical resolutions and deadtimes of common charged particle detectors. Revised November 2011.

Detector Type	Intrinsic Spatial Resolution (rms)	Time Resolution	Dead Time
Resistive plate chamber	$\lesssim 10$ mm	1 ns (50 ps <sup>a</sup> )	—
Streamer chamber	$300\text{ }\mu\text{m}^b$	$2\text{ }\mu\text{s}$	100 ms
Liquid argon drift [7]	$\sim 175\text{--}450\text{ }\mu\text{m}$	$\sim 200$ ns	$\sim 2\text{ }\mu\text{s}$
Scintillation tracker	$\sim 100\text{ }\mu\text{m}$	$100\text{ ps}/n^c$	10 ns
Bubble chamber	$10\text{--}150\text{ }\mu\text{m}$	1 ms	$50\text{ ms}^d$
Proportional chamber	$50\text{--}100\text{ }\mu\text{m}^e$	2 ns	20-200 ns
Drift chamber	$50\text{--}100\text{ }\mu\text{m}$	$2\text{ ns}^f$	20-100 ns
Micro-pattern gas detectors	$30\text{--}40\text{ }\mu\text{m}$	$< 10$ ns	10-100 ns
Silicon strip	pitch/(3 to 7) <sup>g</sup>	few ns <sup>h</sup>	$\lesssim 50\text{ ns}^h$
Silicon pixel	$\lesssim 10\text{ }\mu\text{m}$	few ns <sup>h</sup>	$\lesssim 50\text{ ns}^h$
Emulsion	$1\text{ }\mu\text{m}$	—	—

<sup>a</sup> For multiple-gap RPCs.

<sup>b</sup>  $300\text{ }\mu\text{m}$  is for 1 mm pitch (wirespacing/ $\sqrt{12}$ ).

<sup>c</sup>  $n$  = index of refraction.

<sup>d</sup> Multiple pulsing time.

<sup>e</sup> Delay line cathode readout can give  $\pm 150\text{ }\mu\text{m}$  parallel to anode wire.

<sup>f</sup> For two chambers.

<sup>g</sup> The highest resolution (“7”) is obtained for small-pitch detectors ( $\lesssim 25\text{ }\mu\text{m}$ ) with pulse-height-weighted center finding.

<sup>h</sup> Limited by the readout electronics [8].



### 34.2. Photon detectors

Updated August 2011 by D. Chakraborty (Northern Illinois U) and T. Sumiyoshi (Tokyo Metro U).

Most detectors in high-energy, nuclear, and astrophysics rely on the detection of photons in or near the visible range,  $100 \text{ nm} \lesssim \lambda \lesssim 1000 \text{ nm}$ , or  $E \approx$  a few eV. This range covers scintillation and Cherenkov radiation as well as the light detected in many astronomical observations.

Generally, photodetection involves generating a detectable electrical signal proportional to the (usually very small) number of incident photons. The process involves three distinct steps:

1. generation of a primary photoelectron or electron-hole ( $e$ - $h$ ) pair by an incident photon by the photoelectric or photoconductive effect,
2. amplification of the p.e. signal to detectable levels by one or more multiplicative bombardment steps and/or an avalanche process (usually), and,
3. collection of the secondary electrons to form the electrical signal.

The important characteristics of a photodetector include the following in statistical averages:

1. quantum efficiency (QE or  $\epsilon_Q$ ): the number of primary photoelectrons generated per incident photon ( $0 \leq \epsilon_Q \leq 1$ ; in silicon more than one  $e$ - $h$  pair per incident photon can be generated for  $\lambda \lesssim 165 \text{ nm}$ ),
2. collection efficiency (CE or  $\epsilon_C$ ): the overall acceptance factor other than the generation of photoelectrons ( $0 \leq \epsilon_C \leq 1$ ),
3. gain ( $G$ ): the number of electrons collected for each photoelectron generated,
4. dark current or dark noise: the electrical signal when there is no photon,
5. energy resolution: electronic noise (ENC or  $N_e$ ) and statistical fluctuations in the amplification process compound the Poisson distribution of  $n_\gamma$  photons from a given source:

$$\frac{\sigma(E)}{\langle E \rangle} = \sqrt{\frac{f_N}{n_\gamma \epsilon_Q \epsilon_C} + \left( \frac{N_e}{G n_\gamma \epsilon_Q \epsilon_C} \right)^2}, \quad (34.1)$$

where  $f_N$ , or the excess noise factor (ENF), is the contribution to the energy distribution variance due to amplification statistics [9],

6. dynamic range: the maximum signal available from the detector (this is usually expressed in units of the response to noise-equivalent power, or NEP, which is the optical input power that produces a signal-to-noise ratio of 1),
7. time dependence of the response: this includes the transit time, which is the time between the arrival of the photon and the electrical pulse, and the transit time spread, which contributes to the pulse rise time and width, and
8. rate capability: inversely proportional to the time needed, after the arrival of one photon, to get ready to receive the next.

**Table 34.2:** Representative characteristics of some photodetectors commonly used in particle physics. The time resolution of the devices listed here vary in the 10–2000 ps range.

Type	$\lambda$ (nm)	$\epsilon_Q \epsilon_C$	Gain	Risetime (ns)	Area (mm <sup>2</sup> )	1-p.e noise (Hz)	HV (V)	Price (USD)
PMT*	115–1700	0.15–0.25	$10^3$ – $10^7$	0.7–10	$10^2$ – $10^5$	$10$ – $10^4$	500–3000	100–5000
MCP*	100–650	0.01–0.10	$10^3$ – $10^7$	0.15–0.3	$10^2$ – $10^4$	0.1–200	500–3500	10–6000
HPD*	115–850	0.1–0.3	$10^3$ – $10^4$	7	$10^2$ – $10^5$	$10$ – $10^3$	$\sim 2 \times 10^4$	$\sim 600$
GPM*	115–500	0.15–0.3	$10^3$ – $10^6$	$O(0.1)$	$O(10)$	$10$ – $10^3$	300–2000	$O(10)$
APD	300–1700	$\sim 0.7$	$10$ – $10^8$	$O(1)$	$10$ – $10^3$	$1$ – $10^3$	400–1400	$O(100)$
PPD	320–900	0.15–0.3	$10^5$ – $10^6$	$\sim 1$	1–10	$O(10^6)$	30–60	$O(100)$
VLPC	500–600	$\sim 0.9$	$\sim 5 \times 10^4$	$\sim 10$	1	$O(10^4)$	$\sim 7$	$\sim 1$

\*These devices often come in multi-anode configurations. In such cases, area, noise, and price are to be considered on a “per readout-channel” basis.

The QE is a strong function of the photon wavelength ( $\lambda$ ), and is usually quoted at maximum, together with a range of  $\lambda$  where the QE is comparable to its maximum. Spatial uniformity and linearity with respect to the number of photons are highly desirable in a photodetector’s response.

Optimization of these factors involves many trade-offs and vary widely between applications. For example, while a large gain is desirable, attempts to increase the gain for a given device also increases the ENF and after-pulsing (“echos” of the main pulse). In solid-state devices, a higher QE often requires a compromise in the timing properties. In other types, coverage of large areas by focusing increases the transit time spread.

Other important considerations also are highly application-specific. These include the photon flux and wavelength range, the total area to be covered and the efficiency required, the volume available to accommodate the detectors, characteristics of the environment such as chemical composition, temperature, magnetic field, ambient background, as well as ambient radiation of different types and, mode of operation (continuous or triggered), bias (high-voltage) requirements, power consumption, calibration needs, aging, cost, and so on. Several technologies employing different phenomena for the three steps described above, and many variants within each, offer a wide range of solutions to choose from. The salient features of the main technologies and the common variants are described below. Some key characteristics are summarized in Table 34.2.

**34.2.1. Vacuum photodetectors :** Vacuum photodetectors can be broadly subdivided into three types: photomultiplier tubes, microchannel plates, and hybrid photodetectors.

**34.2.1.1. Photomultiplier tubes:** A versatile class of photon detectors, vacuum photomultiplier tubes (PMT) has been employed by a vast majority of all particle physics experiments to date [9]. Both “transmission-” and “reflection-type” PMT’s are widely used. In the former, the photocathode material is deposited on the inside of a transparent window through which the photons enter, while in the latter, the photocathode material rests on a separate surface that the incident photons strike. The cathode material has a low work function, chosen for the wavelength band of interest. When a photon hits the cathode and liberates an electron (the photoelectric effect), the latter is accelerated and guided by electric fields to impinge on a secondary-emission electrode, or dynode, which then emits a few ( $\sim 5$ ) secondary electrons. The multiplication process is repeated typically 10 times in series to generate a sufficient number of electrons, which are collected at the anode for delivery to the external circuit. The total gain of a PMT depends on the applied high voltage  $V$  as  $G = AV^{kn}$ , where  $k \approx 0.7$ – $0.8$  (depending on the dynode material),  $n$  is the number of dynodes in the chain, and  $A$  a constant (which also depends on  $n$ ). Typically,  $G$  is in the range of  $10^5$ – $10^6$ . Pulse risetimes are usually in the few nanosecond range. With *e.g.* two-level discrimination the effective time resolution can be much better.

A large variety of PMT's, including many just recently developed, covers a wide span of wavelength ranges from infrared (IR) to extreme ultraviolet (XUV) [10]. They are categorized by the window materials, photocathode materials, dynode structures, anode configurations, *etc.* Common window materials are borosilicate glass for IR to near-UV, fused quartz and sapphire ( $\text{Al}_2\text{O}_3$ ) for UV, and  $\text{MgF}_2$  or LiF for XUV. The choice of photocathode materials include a variety of mostly Cs- and/or Sb-based compounds such as CsI, CsTe, bi-alkali (SbRbCs, SbKCs), multi-alkali (SbNa<sub>2</sub>KCs), GaAs(Cs), GaAsP, *etc.* Sensitive wavelengths and peak quantum efficiencies for these materials are summarized in Table 34.3. Typical dynode structures used in PMT's are circular cage, line focusing, box and grid, venetian blind, and fine mesh. In some cases, limited spatial resolution can be obtained by using a mosaic of multiple anodes. Fast PMT's with very large windows—measuring up to 508 mm across—have been developed in recent years for detection of Cherenkov radiation in neutrino experiments such as Super-Kamiokande and KamLAND among many others. Specially prepared low-radioactivity glass is used to make these PMT's, and they are also able to withstand the high pressure of the surrounding liquid.

PMT's are vulnerable to magnetic fields—sometimes even the geomagnetic field causes large orientation-dependent gain changes. A high-permeability metal shield is often necessary. However, proximity-focused PMT's, *e.g.* the fine-mesh types, can be used even in a high magnetic field ( $\geq 1$  T) if the electron drift direction is parallel to the field. CMS uses custom-made vacuum phototriodes (VPT) mounted on the back face of projective lead tungstate crystals to detect scintillation light in the endcap sections of its electromagnetic calorimeters, which are inside a 3.8 T superconducting solenoid. A VPT employs a single dynode (thus,  $G \approx 10$ ) placed close to the photocathode, and a mesh anode plane between the two, to help it cope with the strong magnetic field, which is not too unfavorably oriented with respect to the photodetector axis in the endcaps (within  $25^\circ$ ), but where the radiation level is too high for Avalanche Photodiodes (APD's) like those used in the barrel section.

**34.2.1.2. Microchannel plates:** A typical Microchannel plate (MCP) photodetector consists of one or more  $\sim 2$  mm thick glass plates with densely packed  $O(10 \mu\text{m})$ -diameter cylindrical holes, or “channels”, sitting between the transmission-type photocathode and anode planes, separated by  $O(1 \text{ mm})$  gaps. Instead of discrete dynodes, the inner surface of each cylindrical tube serves as a continuous dynode for the entire cascade of multiplicative bombardments initiated by a photoelectron. Gain fluctuations can be minimized by operating in a saturation mode, whence each channel is only capable of a binary output, but the sum of all channel outputs remains proportional to the number of photons received so long as the photon flux is low enough to ensure that the probability of a single channel receiving more than one photon during a single time gate is negligible. MCP's are thin, offer good spatial resolution, have excellent time resolution ( $\sim 20$  ps), and can tolerate random magnetic fields up to 0.1 T and axial fields up to  $\sim 1$  T. However, they suffer from relatively long recovery time per channel and short lifetime. MCP's are widely employed as image-intensifiers, although not so much in HEP or astrophysics.

**34.2.1.3. Hybrid photon detectors:** Hybrid photon detectors (HPD) combine the sensitivity of a vacuum PMT with the excellent spatial and energy resolutions of a Si sensor [11]. A single photoelectron ejected from the photocathode is accelerated through a potential difference of  $\sim 20$  kV before it impinges on the silicon sensor/anode. The gain nearly equals the maximum number of  $e$ - $h$  pairs that could be created from the entire kinetic energy of the accelerated electron:  $G \approx eV/w$ , where  $e$  is the electronic charge,  $V$  is the applied potential difference, and  $w \approx 3.7$  eV is the mean energy required to create an  $e$ - $h$  pair in Si at room temperature. Since the gain is achieved in a single step, one might expect to have the excellent resolution of a simple Poisson statistic with large mean, but in fact it is even better, thanks to the Fano effect discussed in Sec. 34.7.

Low-noise electronics must be used to read out HPD's if one intends to take advantage of the low fluctuations in gain, *e.g.* when counting small numbers of photons. HPD's can have the same  $\epsilon_Q \epsilon_C$  and window geometries as PMT's and can be segmented down to  $\sim 50$

$\mu\text{m}$ . However, they require rather high biases and will not function in a magnetic field. The exception is proximity-focused devices ( $\Rightarrow$  no (de)magnification) in an axial field. With time resolutions of  $\sim 10$  ps and superior rate capability, proximity-focused HPD's can be an alternative to MCP's. Current applications of HPD's include the CMS hadronic calorimeter and the RICH detector in LHCb. Large-size HPD's with sophisticated focusing may be suitable for future water Cherenkov experiments.

Hybrid APD's (HAPD's) add an avalanche multiplication step following the electron bombardment to boost the gain by a factor of  $\sim 50$ . This affords a higher gain and/or lower electrical bias, but also degrades the signal definition.

**Table 34.3:** Properties of photocathode and window materials commonly used in vacuum photodetectors [10].

Photocathode material	$\lambda$ (nm)	Window material	Peak $\epsilon_Q$ ( $\lambda/\text{nm}$ )
CsI	115–200	$\text{MgF}_2$	0.11 (140)
CsTe	115–320	$\text{MgF}_2$	0.14 (240)
Bi-alkali	300–650	Borosilicate	0.27 (390)
	160–650	Synthetic Silica	0.27 (390)
“Ultra Bi-alkali”	300–650	Borosilicate	0.43 (350)
	160–650	Synthetic Silica	0.43 (350)
Multi-alkali	300–850	Borosilicate	0.20 (360)
	160–850	Synthetic Silica	0.20 (360)
GaAs(Cs)*	160–930	Synthetic Silica	0.23 (280)
GaAsP(Cs)	300–750	Borosilicate	0.50 (500)
InP/InGaAsP <sup>†</sup>	350–1700	Borosilicate	0.01 (1100)

\*Reflection type photocathode is used. <sup>†</sup>Requires cooling to  $\sim -80^\circ\text{C}$ .

**34.2.2. Gaseous photon detectors:** In gaseous photomultipliers (GPM) a photoelectron in a suitable gas mixture initiates an avalanche in a high-field region, producing a large number of secondary impact-ionization electrons. In principle the charge multiplication and collection processes are identical to those employed in gaseous tracking detectors such as multiwire proportional chambers, micromesh gaseous detectors (Micromegas), or gas electron multipliers (GEM). These are discussed in Sec. 34.6.4.

The devices can be divided into two types depending on the photocathode material. One type uses solid photocathode materials much in the same way as PMT's. Since it is resistant to gas mixtures typically used in tracking chambers, CsI is a common choice. In the other type, photoionization occurs on suitable molecules vaporized and mixed in the drift volume. Most gases have photoionization work functions in excess of 10 eV, which would limit their sensitivity to wavelengths far too short. However, vapors of TMAE (tetrakis dimethyl-amine ethylene) or TEA (tri-ethyl-amine), which have smaller work functions (5.3 eV for TMAE and 7.5 eV for TEA), are suited for XUV photon detection [12]. Since devices like GEM's offer sub-mm spatial resolution, GPM's are often used as position-sensitive photon detectors. They can be made into flat panels to cover large areas ( $O(1 \text{ m}^2)$ ), can operate in high magnetic fields, and are relatively inexpensive. Many of the ring imaging Cherenkov (RICH) detectors to date have used GPM's for the detection of Cherenkov light [13]. Special care must be taken to suppress the photon-feedback process in GPM's. It is also important to maintain high purity of the gas as minute traces of  $\text{O}_2$  can significantly degrade the detection efficiency.

**34.2.3. Solid-state photon detectors :** In a phase of rapid development, solid-state photodetectors are competing with vacuum- or gas-based devices for many existing applications and making way for a multitude of new ones. Compared to traditional vacuum- and gaseous photodetectors, solid-state devices are more compact, lightweight, rugged, tolerant to magnetic fields, and often cheaper. They also allow fine pixelization, are easy to integrate into large systems, and can operate at low electric potentials, while matching or exceeding most performance criteria. They are particularly well suited for detection of  $\gamma$ - and X-rays. Except for applications where coverage of very large areas or dynamic range is required, solid-state detectors are proving to be the better choice. Some hybrid devices attempt to combine the best features of different technologies while applications of nanotechnology are opening up exciting new possibilities.

Silicon photodiodes (PD) are widely used in high-energy physics as particle detectors and in a great number of applications (including solar cells!) as light detectors. The structure is discussed in some detail in Sec. 34.7. In its simplest form, the PD is a reverse-biased  $p$ - $n$  junction. Photons with energies above the indirect bandgap energy (wavelengths shorter than about 1050 nm, depending on the temperature) can create  $e$ - $h$  pairs (the photoconductive effect), which are collected on the  $p$  and  $n$  sides, respectively. Often, as in the PD's used for crystal scintillator readout in CLEO, L3, Belle, BaBar, and GLAST, intrinsic silicon is doped to create a  $p$ - $i$ - $n$  structure. The reverse bias increases the thickness of the depleted region; in the case of these particular detectors, to full depletion at a depth of about 100  $\mu\text{m}$ . Increasing the depletion depth decreases the capacitance (and hence electronic noise) and extends the red response. Quantum efficiency can exceed 90%, but falls toward the red because of the increasing absorption length of light in silicon. The absorption length reaches 100  $\mu\text{m}$  at 985 nm. However, since  $G = 1$ , amplification is necessary. Optimal low-noise amplifiers are slow, but, even so, noise limits the minimum detectable signal in room-temperature devices to several hundred photons.

Very large arrays containing  $O(10^7)$  of  $O(10 \mu\text{m}^2)$ -sized photodiodes pixelizing a plane are widely used to photograph all sorts of things from everyday subjects at visible wavelengths to crystal structures with X-rays and astronomical objects from infrared to UV. To limit the number of readout channels, these are made into charge-coupled devices (CCD), where pixel-to-pixel signal transfer takes place over thousands of synchronous cycles with sequential output through shift registers [14]. Thus, high spatial resolution is achieved at the expense of speed and timing precision. Custom-made CCD's have virtually replaced photographic plates and other imagers for astronomy and in spacecraft. Typical QE's exceed 90% over much of the visible spectrum, and "thick" CCD's have useful QE up to  $\lambda = 1 \mu\text{m}$ . Active Pixel Sensor (APS) arrays with a preamplifier on each pixel and CMOS processing afford higher speeds, but are challenged at longer wavelengths. Much R&D is underway to overcome the limitations of both CCD and CMOS imagers.

In APD's, an exponential cascade of impact ionizations initiated by the original photogenerated  $e$ - $h$  pair under a large reverse-bias voltage leads to an avalanche breakdown [15]. As a result, detectable electrical response can be obtained from low-intensity optical signals down to single photons. Excellent junction uniformity is critical, and a guard ring is generally used as a protection against edge breakdown. Well-designed APD's, such as those used in CMS' crystal-based electromagnetic calorimeter, have achieved  $\epsilon_Q \epsilon_C \approx 0.7$  with sub-ns response time. The sensitive wavelength window and gain depend on the semiconductor used. The gain is typically 10–200 in linear and up to  $10^8$  in Geiger mode of operation. Stability and close monitoring of the operating temperature are important for linear-mode operation, and substantial cooling is often necessary. Position-sensitive APD's use time information at multiple anodes to calculate the hit position.

One of the most promising recent developments in the field is that of devices consisting of large arrays ( $O(10^3)$ ) of tiny APD's packed over a small area ( $O(1 \text{ mm}^2)$ ) and operated in a limited Geiger mode [16]. Among different names used for this class of photodetectors, "PPD" (for "Pixelized Photon Detector") is most widely accepted (formerly "SIPM"). Although each cell only offers a binary output, linearity with respect to the number of photons is achieved by summing the

cell outputs in the same way as with a MCP in saturation mode (see above). PPD's are being adopted as the preferred solution for various purposes including medical imaging, *e.g.* positron emission tomography (PET). These compact, rugged, and economical devices allow auto-calibration through decent separation of photoelectron peaks and offer gains of  $O(10^6)$  at a moderate bias voltage ( $\sim 50 \text{ V}$ ). However, the single-photoelectron noise of a PPD, being the logical "or" of  $O(10^3)$  Geiger APD's, is rather large:  $O(1 \text{ MHz/mm}^2)$  at room temperature. PPD's are particularly well-suited for applications where triggered pulses of several photons are expected over a small area, *e.g.* fiber-guided scintillation light. Intense R&D is expected to lower the noise level and improve radiation hardness, resulting in coverage of larger areas and wider applications. Attempts are being made to combine the fabrication of the sensors and the front-end electronics (ASIC) in the same process with the goal of making PPD's and other finely pixelized solid-state photodetectors extremely easy to use.

Of late, much R&D has been directed to  $p$ - $i$ - $n$  diode arrays based on thin polycrystalline diamond films formed by chemical vapor deposition (CVD) on a hot substrate ( $\sim 1000 \text{ K}$ ) from a hydrocarbon-containing gas mixture under low pressure ( $\sim 100 \text{ mbar}$ ). These devices have maximum sensitivity in the extreme- to moderate-UV region [17]. Many desirable characteristics, including high tolerance to radiation and temperature fluctuations, low dark noise, blindness to most of the solar radiation spectrum, and relatively low cost make them ideal for space-based UV/XUV astronomy, measurement of synchrotron radiation, and luminosity monitoring at (future) lepton collider(s).

Visible-light photon counters (VLPC) utilize the formation of an impurity band only 50 meV below the conduction band in As-doped Si to generate strong ( $G \approx 5 \times 10^4$ ) yet sharp response to single photons with  $\epsilon_Q \approx 0.9$  [18]. The smallness of the band gap considerably reduces the gain dispersion. Only a very small bias ( $\sim 7 \text{ V}$ ) is needed, but high sensitivity to infrared photons requires cooling below 10 K. The dark noise increases sharply and exponentially with both temperature and bias. The Run 2 DØ detector used 86000 VLPC's to read the optical signal from its scintillating-fiber tracker and scintillator-strip preshower detectors.

### 34.3. Organic scintillators

Revised August 2011 by Kurtis F. Johnson (FSU).

Organic scintillators are broadly classed into three types, crystalline, liquid, and plastic, all of which utilize the ionization produced by charged particles (see Sec. 33.2 of this *Review*) to generate optical photons, usually in the blue to green wavelength regions [19]. Plastic scintillators are by far the most widely used, liquid organic scintillator is finding increased use, and crystal organic scintillators are practically unused in high-energy physics. Plastic scintillator densities range from 1.03 to 1.20  $\text{g cm}^{-3}$ . Typical photon yields are about 1 photon per 100 eV of energy deposit [20]. A one-cm-thick scintillator traversed by a minimum-ionizing particle will therefore yield  $\approx 2 \times 10^4$  photons. The resulting photoelectron signal will depend on the collection and transport efficiency of the optical package and the quantum efficiency of the photodetector.

Organic scintillator does not respond linearly to the ionization density. Very dense ionization columns emit less light than expected on the basis of  $dE/dx$  for minimum-ionizing particles. A widely used semi-empirical model by Birks posits that recombination and quenching effects between the excited molecules reduce the light yield [21]. These effects are more pronounced the greater the density of the excited molecules. Birks' formula is

$$\frac{d\mathcal{L}}{dx} = \mathcal{L}_0 \frac{dE/dx}{1 + k_B dE/dx}, \quad (34.2)$$

where  $\mathcal{L}$  is the luminescence,  $\mathcal{L}_0$  is the luminescence at low specific ionization density, and  $k_B$  is Birks' constant, which must be determined for each scintillator by measurement. Decay times are in the ns range; rise times are much faster. The high light yield and fast response time allow the possibility of sub-ns timing resolution [22].

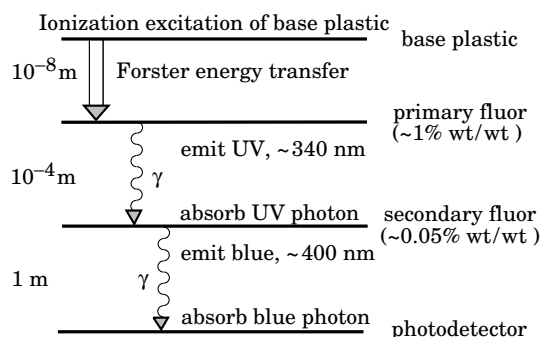
The fraction of light emitted during the decay “tail” can depend on the exciting particle. This allows pulse shape discrimination as a technique to carry out particle identification. Because of the hydrogen content (carbon to hydrogen ratio  $\approx 1$ ) plastic scintillator is sensitive to proton recoils from neutrons. Ease of fabrication into desired shapes and low cost has made plastic scintillator a common detector element. In the form of scintillating fiber it has found widespread use in tracking and calorimetry [23].

Demand for large volume detectors has lead to increased use of liquid organic scintillator, which has the same scintillation mechanism as plastic scintillator, due to its cost advantage. The containment vessel defines the detector shape; photodetectors or waveshifters may be immersed in the liquid.

### 34.3.1. Scintillation mechanism :

A charged particle traversing matter leaves behind it a wake of excited molecules. Certain types of molecules, however, will release a small fraction ( $\approx 3\%$ ) of this energy as optical photons. This process, scintillation, is especially marked in those organic substances which contain aromatic rings, such as polystyrene (PS) and polyvinyltoluene (PVT). Liquids which scintillate include toluene, xylene and pseudocumene.

In fluorescence, the initial excitation takes place via the absorption of a photon, and de-excitation by emission of a longer wavelength photon. Fluors are used as “waveshifters” to shift scintillation light to a more convenient wavelength. Occurring in complex molecules, the absorption and emission are spread out over a wide band of photon energies, and have some overlap, that is, there is some fraction of the emitted light which can be re-absorbed [24]. This “self-absorption” is undesirable for detector applications because it causes a shortened attenuation length. The wavelength difference between the major absorption and emission peaks is called the Stokes’ shift. It is usually the case that the greater the Stokes’ shift, the smaller the self absorption thus, a large Stokes’ shift is a desirable property for a fluor.



**Figure 34.1:** Cartoon of scintillation “ladder” depicting the operating mechanism of organic scintillator. Approximate fluor concentrations and energy transfer distances for the separate sub-processes are shown.

The plastic scintillators used in high-energy physics are binary or ternary solutions of selected fluors in a plastic base containing aromatic rings. (See the appendix in Ref. 25 for a comprehensive list of components.) Virtually all plastic scintillators contain as a base either PVT or PS. PVT-based scintillator can be up to 50% brighter.

Ionization in the plastic base produces UV photons with short attenuation length (several mm). Longer attenuation lengths are obtained by dissolving a “primary” fluor in high concentration (1% by weight) into the base, which is selected to efficiently re-radiate absorbed energy at wavelengths where the base is more transparent (see Fig. 34.1).

The primary fluor has a second important function. The decay time of the scintillator base material can be quite long – in pure polystyrene it is 16 ns, for example. The addition of the primary fluor in high concentration can shorten the decay time by an order of magnitude and increase the total light yield. At the concentrations used (1% and

greater), the average distance between a fluor molecule and an excited base unit is around 100 Å, much less than a wavelength of light. At these distances the predominant mode of energy transfer from base to fluor is not the radiation of a photon, but a resonant dipole-dipole interaction, first described by Foerster, which strongly couples the base and fluor [26]. The strong coupling sharply increases the speed and the light yield of the plastic scintillators.

Unfortunately, a fluor which fulfills other requirements is usually not completely adequate with respect to emission wavelength or attenuation length, so it is necessary to add yet another waveshifter (the “secondary” fluor), at fractional percent levels, and occasionally a third (not shown in Fig. 34.1).

External wavelength shifters are widely used to aid light collection in complex geometries. Scintillation light is captured by a lightpipe comprising a wave-shifting fluor dissolved in a nonscintillating base. The wavelength shifter must be insensitive to ionizing radiation and Cherenkov light. A typical wavelength shifter uses an acrylic base because of its good optical qualities, a single fluor to shift the light emerging from the plastic scintillator to the blue-green, and contains ultra-violet absorbing additives to deaden response to Cherenkov light.

### 34.3.2. Caveats and cautions :

Plastic scintillators are reliable, robust, and convenient. However, they possess quirks to which the experimenter must be alert. Exposure to solvent vapors, high temperatures, mechanical flexing, irradiation, or rough handling will aggravate the process. A particularly fragile region is the surface which can “craze” develop microcracks which degrade its transmission of light by total internal reflection. Crazing is particularly likely where oils, solvents, or *fingerprints* have contacted the surface.

They have a long-lived luminescence which does not follow a simple exponential decay. Intensities at the  $10^{-4}$  level of the initial fluorescence can persist for hundreds of ns [19,27].

They will decrease their light yield with increasing partial pressure of oxygen. This can be a 10% effect in an artificial atmosphere [28]. It is not excluded that other gases may have similar quenching effects.

Their light yield may be changed by a magnetic field. The effect is very nonlinear and apparently not all types of plastic scintillators are so affected. Increases of  $\approx 3\%$  at 0.45 T have been reported [29]. Data are sketchy and mechanisms are not understood.

Irradiation of plastic scintillators creates color centers which absorb light more strongly in the UV and blue than at longer wavelengths. This poorly understood effect appears as a reduction both of light yield and attenuation length. Radiation damage depends not only on the integrated dose, but on the dose rate, atmosphere, and temperature, before, during and after irradiation, as well as the materials properties of the base such as glass transition temperature, polymer chain length, *etc.* Annealing also occurs, accelerated by the diffusion of atmospheric oxygen and elevated temperatures. The phenomena are complex, unpredictable, and not well understood [30]. Since color centers are less disruptive at longer wavelengths, the most reliable method of mitigating radiation damage is to shift emissions at every step to the longest practical wavelengths, *e.g.*, utilize fluors with large Stokes’ shifts (aka the “Better red than dead” strategy).

### 34.3.3. Scintillating and wavelength-shifting fibers :

The clad optical fiber comprising scintillator and wavelength shifter (WLS) is particularly useful [31]. Since the initial demonstration of the scintillating fiber (SCIFI) calorimeter [32], SCIFI techniques have become mainstream [33]. SCIFI calorimeters are fast, dense, radiation hard, and can have leadglass-like resolution. SCIFI trackers can handle high rates and are radiation tolerant, but the low photon yield at the end of a long fiber (see below) forces the use of sensitive photodetectors. WLS scintillator readout of a calorimeter allows a very high level of hermeticity since the solid angle blocked by the fiber on its way to the photodetector is very small. The sensitive region of scintillating fibers can be controlled by splicing them onto clear (non-scintillating/non-WLS) fibers.

A typical configuration would be fibers with a core of polystyrene-based scintillator or WLS (index of refraction  $n = 1.59$ ), surrounded

by a cladding of PMMA ( $n = 1.49$ ) a few microns thick, or, for added light capture, with another cladding of fluorinated PMMA with  $n = 1.42$ , for an overall diameter of 0.5 to 1 mm. The fiber is drawn from a boule and great care is taken during production to ensure that the intersurface between the core and the cladding has the highest possible uniformity and quality, so that the signal transmission via total internal reflection has a low loss. The fraction of generated light which is transported down the optical pipe is denoted the capture fraction and is about 6% for the single-clad fiber and 10% for the double-clad fiber. The number of photons from the fiber available at the photodetector is always smaller than desired, and increasing the light yield has proven difficult. A minimum-ionizing particle traversing a high-quality 1 mm diameter fiber perpendicular to its axis will produce fewer than 2000 photons, of which about 200 are captured. Attenuation may eliminate 95% of these photons in a large collider tracker.

A scintillating or WLS fiber is often characterized by its attenuation length, over which the signal is attenuated to  $1/e$  of its original value. Many factors determine the attenuation length, including the importance of re-absorption of emitted photons by the polymer base or dissolved fluors, the level of crystallinity of the base polymer, and the quality of the total internal reflection boundary [34]. Attenuation lengths of several meters are obtained by high quality fibers. However, it should be understood that the attenuation length is not the sole measure of fiber quality. Among other things, it is not constant with distance from the excitation source and it is wavelength dependent.

### 34.4. Inorganic scintillators

Revised November 2015 by R.-Y. Zhu (California Institute of Technology) and C.L. Woody (BNL).

Inorganic crystals form a class of scintillating materials with much higher densities than organic plastic scintillators (typically  $\sim 4\text{--}8\text{ g/cm}^3$ ) with a variety of different properties for use as scintillation detectors. Due to their high density and high effective atomic number, they can be used in applications where high stopping power or a high conversion efficiency for electrons or photons is required. These include total absorption electromagnetic calorimeters (see Sec. 34.9.1), which consist of a totally active absorber (as opposed to a sampling calorimeter), as well as serving as gamma ray detectors over a wide range of energies. Many of these crystals also have very high light output, and can therefore provide excellent energy resolution down to very low energies ( $\sim$  few hundred keV).

Some crystals are intrinsic scintillators in which the luminescence is produced by a part of the crystal lattice itself. However, other crystals require the addition of a dopant, typically fluorescent ions such as thallium (Tl) or cerium (Ce) which is responsible for producing the scintillation light. However, in both cases, the scintillation mechanism is the same. Energy is deposited in the crystal by ionization, either directly by charged particles, or by the conversion of photons into electrons or positrons which subsequently produce ionization. This energy is transferred to the luminescent centers which then radiate scintillation photons. The light yield  $L$  in terms of the number of scintillation photons produced per MeV of energy deposit in the crystal can be expressed as [35]

$$L = 10^6 \cdot S \cdot Q / (\beta \cdot E_g), \quad (34.3)$$

where  $\beta \cdot E_g$  is the energy required to create an e-h pair expressed as a multiple of the band gap energy  $E_g$  (eV),  $S$  is the efficiency of energy transfer to the luminescent center and  $Q$  is the quantum efficiency of the luminescent center. The values of  $\beta$ ,  $S$  and  $Q$  are crystal dependent and are the main factors in determining the intrinsic light yield of the scintillator. The decay time of the scintillator is mainly dominated by the decay time of the luminescent center.

Table 34.4 lists the basic properties of some commonly used inorganic crystals. NaI(Tl) is one of the most common and widely used scintillators, with an emission that is well matched to a bialkali photomultiplier tube, but it is highly hygroscopic and difficult to work with, and has a rather low density. CsI(Tl) and CsI(Na) have high light yield, low cost, and are mechanically robust (high plasticity and

resistance to cracking). However, they need careful surface treatment and are slightly and highly hygroscopic respectively. Pure CsI has identical mechanical properties as CsI(Tl), but faster emission at shorter wavelength and a much lower light output. BaF<sub>2</sub> has a fast component with a sub-nanosecond decay time, and is the fastest known scintillator. However, it also has a slow component with a much longer decay time ( $\sim 630$  ns). Bismuth germanate (Bi<sub>4</sub>Ge<sub>3</sub>O<sub>12</sub> or BGO) has a high density, and consequently a short radiation length  $X_0$  and Molière radius  $R_M$ . Similar to CsI(Tl), BGO's emission is well-matched to the spectral sensitivity of photodiodes, and it is easy to handle and not hygroscopic. Lead tungstate (PbWO<sub>4</sub> or PWO) has a very high density, with a very short  $X_0$  and  $R_M$ , but its intrinsic light yield is rather low.

Cerium doped lutetium oxyorthosilicate (Lu<sub>2</sub>SiO<sub>5</sub>:Ce, or LSO:Ce) [36] and cerium doped lutetium-yttrium oxyorthosilicate (Lu<sub>2(1-x)</sub>Y<sub>2x</sub>SiO<sub>5</sub>, LYSO:Ce) [37] are dense crystal scintillators which have a high light yield and a fast decay time. Only the properties of LSO:Ce are listed in Table 34.4 since the properties of LYSO:Ce are similar to that of LSO:Ce except a slightly lower density than LSO:Ce depending on the yttrium fraction in LYSO:Ce. This material is also featured with excellent radiation hardness [38], so is expected to be used where extraordinary radiation hardness is required.

Also listed in Table 34.4 are other fluoride crystals such as PbF<sub>2</sub> as a Cherenkov material and CeF<sub>3</sub>, which have been shown to provide excellent energy resolution in calorimeter applications. Table 34.4 also includes cerium doped lanthanum tri-halides, such as LaBr<sub>3</sub> [39] and CeBr<sub>3</sub> [40], which are brighter and faster than LSO:Ce, but they are highly hygroscopic and have a lower density. The FWHM energy resolution measured for these materials coupled to a PMT with bi-alkali photocathode for 0.662 MeV  $\gamma$ -rays from a <sup>137</sup>Cs source is about 3%, and has recently been improved to 2% by co-doping with cerium and strontium [41], which is the best among all inorganic crystal scintillators. For this reason, LaBr<sub>3</sub> and CeBr<sub>3</sub> are expected to be used in applications where a good energy resolution for low energy photons are required, such as homeland security.

Beside the crystals listed in Table 34.4, a number of new crystals are being developed that may have potential applications in high energy or nuclear physics. Of particular interest is the family of yttrium and lutetium perovskites and garnet, which include YAP (YAlO<sub>3</sub>:Ce), LuAP (LuAlO<sub>3</sub>:Ce), YAG (Y<sub>3</sub>Al<sub>5</sub>O<sub>12</sub>:Ce) and LuAG (Lu<sub>3</sub>Al<sub>5</sub>O<sub>12</sub>:Ce) and their mixed compositions. These have been shown to be linear over a large energy range [42], and have the potential for providing good intrinsic energy resolution.

Aiming at the best jet-mass resolution inorganic scintillators are being investigated for HEP calorimeters with dual readout for both Cherenkov and scintillation light to be used at future linear colliders. These materials may be used for an electromagnetic calorimeter [43] or a homogeneous hadronic calorimetry (HHCAL) detector concept, including both electromagnetic and hadronic parts [44]. Because of the unprecedented volume (70 to 100 m<sup>3</sup>) foreseen for the HHCAL detector concept the materials must be (1) dense (to minimize the leakage) and (2) cost-effective. It should also be UV transparent (for effective collection of the Cherenkov light) and allow for a clear discrimination between the Cherenkov and scintillation light. The preferred scintillation light is thus at a longer wavelength, and not necessarily bright or fast. Dense crystals, scintillating glasses and ceramics offer a very attractive implementation for this detector concept [45].

The fast scintillation light provides timing information about electromagnetic interactions and showers, which may be used to mitigate pile-up effects and/or for particle identification since the time development of electromagnetic and hadronic showers, as well as minimum ionizing particles, are different. The timing information is primarily determined by the scintillator rise time and decay time, and the number of photons produced. For fast timing, it is important to have a large number of photons emitted in the initial part of the scintillation pulse, e.g. in the first ns, since one is often measuring the arrival time of the particle in the crystal using the leading edge of the light pulse. A good example of this is BaF<sub>2</sub>, which has  $\sim 10\%$  of its light in its fast component with a decay time of  $< 1$  ns. The

light propagation can spread out the arrival time of the scintillation photons at the photodetector due to time dispersion [46]. The time response of the photodetector also plays a major role in achieving good time resolution with fast scintillating crystals.

Table 34.4 gives the light output of other crystals relative to NaI(Tl) and their dependence to the temperature variations measured for 1.5  $X_0$  cube crystal samples with a Tyvek paper wrapping and a full end face coupled to a photodetector [47]. The quantum efficiencies of the photodetector is taken out to facilitate a direct comparison of crystal's light output. However, the useful signal produced by a scintillator is usually quoted in terms of the number of photoelectrons per MeV produced by a given photodetector. The relationship between the number of photons/MeV produced ( $L$ ) and photoelectrons/MeV detected ( $N_{p.e.}/\text{MeV}$ ) involves the factors for the light collection efficiency ( $LC$ ) and the quantum efficiency ( $QE$ ) of the photodetector:

$$N_{p.e.}/\text{MeV} = L \cdot LC \cdot QE. \quad (34.4)$$

$LC$  depends on the size and shape of the crystal, and includes effects such as the transmission of scintillation light within the crystal (i.e., the bulk attenuation length of the material), scattering from within the crystal, reflections and scattering from the crystal surfaces, and re-bouncing back into the crystal by wrapping materials. These factors can vary considerably depending on the sample, but can be in the range of  $\sim 10$ –60%. The internal light transmission depends on the intrinsic properties of the material, e.g. the density and type of the scattering centers and defects that can produce internal absorption within the crystal, and can be highly affected by factors such as radiation damage, as discussed below.

The quantum efficiency depends on the type of photodetector used to detect the scintillation light, which is typically  $\sim 15$ –30% for photomultiplier tubes and  $\sim 70\%$  for silicon photodiodes for visible wavelengths. The quantum efficiency of the detector is usually highly wavelength dependent and should be matched to the particular crystal of interest to give the highest quantum yield at the wavelength

corresponding to the peak of the scintillation emission. Fig. 34.2 shows the quantum efficiencies of two photodetectors, a Hamamatsu R2059 PMT with bi-alkali cathode and quartz window and a Hamamatsu S8664 avalanche photodiode (APD) as a function of wavelength. Also shown in the figure are emission spectra of three crystal scintillators, BGO, LSO:Ce/LYSO:Ce and CsI(Tl), and the numerical values of the emission weighted quantum efficiency. The area under each emission spectrum is proportional to crystal's light yield, as shown in Table 34.4, where the quantum efficiencies of the photodetector has been taken out. Results with different photodetectors can be significantly different. For example, the response of CsI(Tl) relative to NaI(Tl) with a standard photomultiplier tube with a bi-alkali photo-cathode, e.g. Hamamatsu R2059, would be 45 rather than 165 because of the photomultiplier's low quantum efficiency at longer wavelengths. For scintillators which emit in the UV, a detector with a quartz window should be used.

For very low energy applications (typically below 1 MeV), non-proportionality of the scintillation light yield may be important. It has been known for a long time that the conversion factor between the energy deposited in a crystal scintillator and the number of photons produced is not constant. It is also known that the energy resolution measured by all crystal scintillators for low energy  $\gamma$ -rays is significantly worse than the contribution from photo-electron statistics alone, indicating an intrinsic contribution from the scintillator itself. Precision measurement using low energy electron beam shows that this non-proportionality is crystal dependent [48]. Recent study on this issue also shows that this effect is also sample dependent even for the same crystal [49]. Further work is therefore needed to fully understand this subject.

One important issue related to the application of a crystal scintillator is its radiation hardness. Stability of its light output, or the ability to track and monitor the variation of its light output in a radiation environment, is required for high resolution and precision calibration [50]. All known crystal scintillators suffer from ionization dose induced radiation damage [51], where a common damage

**Table 34.4:** Properties of several inorganic crystals. Most of the notation is defined in Sec. 6 of this *Review*.

Parameter:	$\rho$	MP	$X_0^*$	$R_M^*$	$dE^*/dx$	$\lambda_I^*$	$\tau_{\text{decay}}$	$\lambda_{\text{max}}$	$n^\ddagger$	Relative output	Hygroscopic?	$d(LY)/dT$
Units:	$\text{g}/\text{cm}^3$	$^\circ\text{C}$	cm	cm	$\text{MeV}/\text{cm}$	cm	ns	nm				$\%/^\circ\text{C}^\ddagger$
NaI(Tl)	3.67	651	2.59	4.13	4.8	42.9	245	410	1.85	100	yes	-0.2
BGO	7.13	1050	1.12	2.23	9.0	22.8	300	480	2.15	21	no	-0.9
BaF <sub>2</sub>	4.89	1280	2.03	3.10	6.5	30.7	650 <sup>s</sup> 0.9 <sup>f</sup>	300 <sup>s</sup> 220 <sup>f</sup>	1.50	36 <sup>s</sup> 4.1 <sup>f</sup>	no	-1.9 <sup>s</sup> 0.1 <sup>f</sup>
CsI(Tl)	4.51	621	1.86	3.57	5.6	39.3	1220	550	1.79	165	slight	0.4
CsI(Na)	4.51	621	1.86	3.57	5.6	39.3	690	420	1.84	88	yes	0.4
CsI(pure)	4.51	621	1.86	3.57	5.6	39.3	30 <sup>s</sup> 6 <sup>f</sup>	310	1.95	3.6 <sup>s</sup> 1.1 <sup>f</sup>	slight	-1.4
PbWO <sub>4</sub>	8.30	1123	0.89	2.00	10.1	20.7	30 <sup>s</sup> 10 <sup>f</sup>	425 <sup>s</sup> 420 <sup>f</sup>	2.20	0.3 <sup>s</sup> 0.077 <sup>f</sup>	no	-2.5
LSO(Ce)	7.40	2050	1.14	2.07	9.6	20.9	40	402	1.82	85	no	-0.2
PbF <sub>2</sub>	7.77	824	0.93	2.21	9.4	21.0	-	-	-	Cherenkov	no	-
CeF <sub>3</sub>	6.16	1460	1.70	2.41	8.42	23.2	30	340	1.62	7.3	no	0
LaBr <sub>3</sub> (Ce)	5.29	783	1.88	2.85	6.90	30.4	20	356	1.9	180	yes	0.2
CeBr <sub>3</sub>	5.23	722	1.96	2.97	6.65	31.5	17	371	1.9	165	yes	-0.1

\* Numerical values calculated using formulae in this review.

$^\ddagger$  Refractive index at the wavelength of the emission maximum.

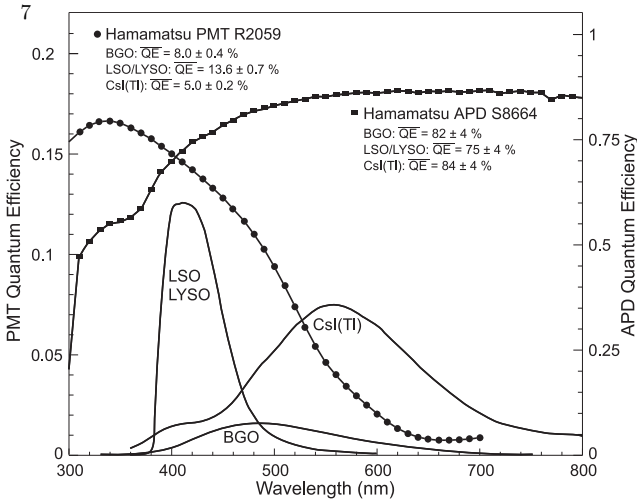
$^\dagger$  Relative light output measured for samples of 1.5  $X_0$  cube with a Tyvek paper wrapping and a full end face coupled to a photodetector. The quantum efficiencies of the photodetector are taken out.

$^\ddagger$  Variation of light yield with temperature evaluated at the room temperature.

$f$  = fast component,  $s$  = slow component



phenomenon is the appearance of radiation induced absorption caused by the formation of color centers originated from the impurities or point defects in the crystal. This radiation induced absorption reduces the light attenuation length in the crystal, and hence its light output. For crystals with high defect density, a severe reduction of light attenuation length may cause a distortion of the light response uniformity, leading to a degradation of the energy resolution. Additional radiation damage effects may include a reduced intrinsic scintillation light yield (damage to the luminescent centers) and an increased phosphorescence (afterglow). For crystals to be used in a high precision calorimeter in a radiation environment, its scintillation mechanism must not be damaged and its light attenuation length in the expected radiation environment must be long enough so that its light response uniformity, and thus its energy resolution, does not change.



**Figure 34.2:** The quantum efficiencies of two photodetectors, a Hamamatsu R2059 PMT with bi-alkali cathode and a Hamamatsu S8664 avalanche photodiode (APD), are shown as a function of wavelength. Also shown in the figure are emission spectra of three crystal scintillators, BGO, LSO and CsI(Tl), and the numerical values of the emission weighted quantum efficiencies. The area under each emission spectrum is proportional to crystal's light yield.

While radiation damage induced by ionization dose is well understood [52], investigation is on-going to understand radiation damage caused by hadrons, including both charged hadrons and neutrons. Two additional fundamental processes may cause defects by hadrons: displacement damage and nuclear breakup. While charged hadrons can produce all three types of damage (and it's often difficult to separate them), neutrons can produce only the last two, and electrons and photons only produce ionization damage. Studies on hadron induced radiation damage to lead tungstate [53] show a proton-specific damage component caused by fragments from fission induced in lead and tungsten by particles in the hadronic shower. The fragments cause a severe, local damage to the crystalline lattice due to their extremely high energy loss over a short distance [53]. Studies on neutron-specific damage in lead tungstate [54] up to  $4 \times 10^{19}$  n/cm<sup>2</sup> show no neutron-specific damage in PWO [55].

Most of the crystals listed in Table 34.4 have been used in high energy or nuclear physics experiments when the ultimate energy resolution for electrons and photons is desired. Examples are the Crystal Ball NaI(Tl) calorimeter at SPEAR, the L3 BGO calorimeter at LEP, the CLEO CsI(Tl) calorimeter at CESR, the KTeV CsI calorimeter at the Tevatron, the BaBar, BELLE and BES II CsI(Tl) calorimeters at PEP-II, KEK and BEPC III. Because of their high density and relative low cost, PWO calorimeters are used by CMS and ALICE at LHC, by CLAS and PrimEx at CEBAF and by PANDA at GSI, and PbF<sub>2</sub> calorimeters are used by the A4 experiment at MAINZ

and by the g-2 experiment at Fermilab. A LYSO:Ce calorimeter is being constructed by the COMET experiment at J-PARC.

### 34.5. Cherenkov detectors

Revised August 2015 by B.N. Ratcliff (SLAC).

Although devices using Cherenkov radiation are often thought of as only particle identification (PID) detectors, in practice they are used over a much broader range of applications including; (1) fast particle counters; (2) hadronic PID; and (3) tracking detectors performing complete event reconstruction. Examples of applications from each category include; (1) the Quartic fast timing counter designed to measure small angle scatters at the LHC [56]; (2) the hadronic PID detectors at the B factory detectors—DIRC in BaBar [57] and the aerogel threshold Cherenkov in Belle [58]; and (3) large water Cherenkov counters such as Super-Kamiokande [59]. Cherenkov counters contain two main elements; (1) a radiator through which the charged particle passes, and (2) a photodetector. As Cherenkov radiation is a weak source of photons, light collection and detection must be as efficient as possible. The refractive index  $n$  and the particle's path length through the radiator  $L$  appear in the Cherenkov relations allowing the tuning of these quantities for particular applications.

Cherenkov detectors utilize one or more of the properties of Cherenkov radiation discussed in the Passages of Particles through Matter section (Sec. 33 of this *Review*): the prompt emission of a light pulse; the existence of a velocity threshold for radiation; and the dependence of the Cherenkov cone half-angle  $\theta_c$  and the number of emitted photons on the velocity of the particle and the refractive index of the medium.

The number of photoelectrons ( $N_{p.e.}$ ) detected in a given device is

$$N_{p.e.} = L \frac{\alpha^2 z^2}{r_e m_e c^2} \int \epsilon(E) \sin^2 \theta_c(E) dE, \quad (34.5)$$

where  $\epsilon(E)$  is the efficiency for collecting the Cherenkov light and transducing it into photoelectrons, and  $\alpha^2/(r_e m_e c^2) = 370 \text{ cm}^{-1} \text{ eV}^{-1}$ .

The quantities  $\epsilon$  and  $\theta_c$  are functions of the photon energy  $E$ . As the typical energy dependent variation of the index of refraction is modest, a quantity called the *Cherenkov detector quality factor*  $N_0$  can be defined as

$$N_0 = \frac{\alpha^2 z^2}{r_e m_e c^2} \int \epsilon dE, \quad (34.6)$$

so that, taking  $z = 1$  (the usual case in high-energy physics),

$$N_{p.e.} \approx L N_0 \langle \sin^2 \theta_c \rangle. \quad (34.7)$$

This definition of the quality factor  $N_0$  is not universal, nor, indeed, very useful for those common situations where  $\epsilon$  factorizes as  $\epsilon = \epsilon_{\text{coll}} \epsilon_{\text{det}}$  with the geometrical photon collection efficiency ( $\epsilon_{\text{coll}}$ ) varying substantially for different tracks while the photon detector efficiency ( $\epsilon_{\text{det}}$ ) remains nearly track independent. In this case, it can be useful to explicitly remove ( $\epsilon_{\text{coll}}$ ) from the definition of  $N_0$ . A typical value of  $N_0$  for a photomultiplier (PMT) detection system working in the visible and near UV, and collecting most of the Cherenkov light, is about  $100 \text{ cm}^{-1}$ . Practical counters, utilizing a variety of different photodetectors, have values ranging between about 30 and  $180 \text{ cm}^{-1}$ . Radiators can be chosen from a variety of transparent materials (Sec. 33 of this *Review* and Table 6.1). In addition to refractive index, the choice requires consideration of factors such as material density, radiation length and radiation hardness, transmission bandwidth, absorption length, chromatic dispersion, optical workability (for solids), availability, and cost. When the momenta of particles to be identified is high, the refractive index must be set close to one, so that the photon yield per unit length is low and a long particle path in the radiator is required. Recently, the gap in refractive index that has traditionally existed between gases and liquid or solid materials has been partially closed with transparent *silica aerogels* with indices that range between about 1.007 and 1.13.

Cherenkov counters may be classified as either *imaging* or *threshold* types, depending on whether they do or do not make use of Cherenkov

angle ( $\theta_c$ ) information. Imaging counters may be used to track particles as well as identify them. The recent development of very fast photodetectors such as micro-channel plate PMTs (MCP PMT) (see Sec. 34.2 of this *Review*) also potentially allows very fast Cherenkov based time of flight (TOF) detectors of either class [60]. The track timing resolution of imaging detectors can be extremely good as it scales approximately as  $\frac{1}{\sqrt{N_{p.e.}}}$ .

*Threshold* Cherenkov detectors [61], in their simplest form, make a yes/no decision based on whether the particle is above or below the Cherenkov threshold velocity  $\beta_t = 1/n$ . A straightforward enhancement of such detectors uses the number of observed photoelectrons (or a calibrated pulse height) to discriminate between species or to set probabilities for each particle species [62]. This strategy can increase the momentum range of particle separation by a modest amount (to a momentum some 20% above the threshold momentum of the heavier particle in a typical case).

Careful designs give  $\langle \epsilon_{\text{coll}} \rangle \gtrsim 90\%$ . For a photomultiplier with a typical alkali cathode,  $\int \epsilon_{\text{det}} dE \approx 0.27 \text{ eV}$ , so that

$$N_{p.e.}/L \approx 90 \text{ cm}^{-1} \langle \sin^2 \theta_c \rangle \quad (\text{i.e., } N_0 = 90 \text{ cm}^{-1}). \quad (34.8)$$

Suppose, for example, that  $n$  is chosen so that the threshold for species  $a$  is  $p_t$ ; that is, at this momentum species  $a$  has velocity  $\beta_a = 1/n$ . A second, lighter, species  $b$  with the same momentum has velocity  $\beta_b$ , so  $\cos \theta_c = \beta_a/\beta_b$ , and

$$N_{p.e.}/L \approx 90 \text{ cm}^{-1} \frac{m_a^2 - m_b^2}{p_t^2 + m_a^2}. \quad (34.9)$$

For  $K/\pi$  separation at  $p = p_t = 1(5) \text{ GeV}/c$ ,  $N_{p.e.}/L \approx 16(0.8) \text{ cm}^{-1}$  for  $\pi$ 's and (by design) 0 for  $K$ 's.

For limited path lengths  $N_{p.e.}$  will usually be small. The overall efficiency of the device is controlled by Poisson fluctuations, which can be especially critical for separation of species where one particle type is dominant. Moreover, the effective number of photoelectrons is often less than the average number calculated above due to additional equivalent noise from the photodetector (see the discussion of the excess noise factor in Sec. 34.2 of this *Review*). It is common to design for at least 10 photoelectrons for the high velocity particle in order to obtain a robust counter. As rejection of the particle that is below threshold depends on *not* seeing a signal, electronic and other background noise, especially overlapping tracks, can be important. Physics sources of light production for the below threshold particle, such as decay to an above threshold particle, scintillation light, or the production of delta rays in the radiator, often limit the separation attainable, and need to be carefully considered. Well designed, modern multi-channel counters, such as the ACC at Belle [58], can attain adequate particle separation performance over a substantial momentum range.

*Imaging* counters make the most powerful use of the information available by measuring the ring-correlated angles of emission of the individual Cherenkov photons. They typically provide positive ID information both for the “wanted” and the “unwanted” particles, thus reducing mis-identification substantially. Since low-energy photon detectors can measure only the position (and, perhaps, a precise detection time) of the individual Cherenkov photons (not the angles directly), the photons must be “imaged” onto a detector so that their angles can be derived [63]. Typically the optics map the Cherenkov cone onto (a portion of) a distorted “circle” at the photodetector. Though the imaging process is directly analogous to familiar imaging techniques used in telescopes and other optical instruments, there is a somewhat bewildering variety of methods used in a wide variety of counter types with different names. Some of the imaging methods used include (1) focusing by a lens or mirror; (2) proximity focusing (i.e., focusing by limiting the emission region of the radiation); and (3) focusing through an aperture (a pinhole). In addition, the prompt Cherenkov emission coupled with the speed of some modern photon detectors allows the use of (4) time imaging, a method which is little used in conventional imaging technology, and may allow some separation with particle TOF. Finally, (5) correlated tracking (and

event reconstruction) can be performed in large water counters by combining the individual space position and time of each photon together with the constraint that Cherenkov photons are emitted from each track at the same polar angle (Sec. 35.3.1 of this *Review*).

In a simple model of an imaging PID counter, the fractional error on the particle velocity ( $\delta_\beta$ ) is given by

$$\delta_\beta = \frac{\sigma_\beta}{\beta} = \tan \theta_c \sigma(\theta_c), \quad (34.10)$$

where

$$\sigma(\theta_c) = \frac{\langle \sigma(\theta_i) \rangle}{\sqrt{N_{p.e.}}} \oplus C, \quad (34.11)$$

and  $\langle \sigma(\theta_i) \rangle$  is the average single photoelectron resolution, as defined by the optics, detector resolution and the intrinsic chromaticity spread of the radiator index of refraction averaged over the photon detection bandwidth.  $C$  combines a number of other contributions to resolution including, (1) correlated terms such as tracking, alignment, and multiple scattering, (2) hit ambiguities, (3) background hits from random sources, and (4) hits coming from other tracks. The actual separation performance is also limited by physics effects such as decays in flight and particle interactions in the material of the detector. In many practical cases, the performance is limited by these effects.

For a  $\beta \approx 1$  particle of momentum ( $p$ ) well above threshold entering a radiator with index of refraction ( $n$ ), the number of  $\sigma$  separation ( $N_\sigma$ ) between particles of mass  $m_1$  and  $m_2$  is approximately

$$N_\sigma \approx \frac{|m_1^2 - m_2^2|}{2p^2 \sigma(\theta_c) \sqrt{n^2 - 1}}. \quad (34.12)$$

In practical counters, the angular resolution term  $\sigma(\theta_c)$  varies between about 0.1 and 5 mrad depending on the size, radiator, and photodetector type of the particular counter. The range of momenta over which a particular counter can separate particle species extends from the point at which the number of photons emitted becomes sufficient for the counter to operate efficiently as a threshold device ( $\sim 20\%$  above the threshold for the lighter species) to the value in the imaging region given by the equation above. For example, for  $\sigma(\theta_c) = 2 \text{ mrad}$ , a fused silica radiator ( $n = 1.474$ ), or a fluorocarbon gas radiator ( $\text{C}_5\text{F}_{12}$ ,  $n = 1.0017$ ), would separate  $\pi/K$ 's from the threshold region starting around  $0.15(3) \text{ GeV}/c$  through the imaging region up to about  $4.2(18) \text{ GeV}/c$  at better than  $3\sigma$ .

Many different imaging counters have been built during the last several decades [60]. Among the earliest examples of this class of counters are the very limited acceptance Differential Cherenkov detectors, designed for particle selection in high momentum beam lines. These devices use optical focusing and/or geometrical masking to select particles having velocities in a specified region. With careful design, a velocity resolution of  $\sigma_\beta/\beta \approx 10^{-4}$ – $10^{-5}$  can be obtained [61].

Practical multi-track Ring-Imaging Cherenkov detectors (generically called RICH counters) are a more recent development. RICH counters are sometimes further classified by ‘generations’ that differ based on historical timing, performance, design, and photodetection techniques.

Prototypical examples of first generation RICH counters are those used in the DELPHI and SLD detectors at the LEP and SLC Z factory  $e^+e^-$  colliders [60]. They have both liquid ( $\text{C}_6\text{F}_{14}$ ,  $n = 1.276$ ) and gas ( $\text{C}_5\text{F}_{12}$ ,  $n = 1.0017$ ) radiators, the former being proximity imaged with the latter using mirrors. The phototransducers are a TPC/wire-chamber combination. They are made sensitive to photons by doping the TPC gas (usually, ethane/methane) with  $\sim 0.05\%$  TMAE (tetrakis(dimethylamino)ethylene). Great attention to detail is required, (1) to avoid absorbing the UV photons to which TMAE is sensitive, (2) to avoid absorbing the single photoelectrons as they drift in the long TPC, and (3) to keep the chemically active TMAE vapor from interacting with materials in the system. In spite of their unforgiving operational characteristics, these counters attained good  $e/\pi/K/p$  separation over wide momentum ranges (from about 0.25 to  $20 \text{ GeV}/c$ ) during several years of operation at LEP and SLC. Related but smaller acceptance devices include the OMEGA RICH



at the CERN SPS, and the RICH in the balloon-borne CAPRICE detector [60].

Later generation counters [60] generally operate at much higher rates, with more detection channels, than the first generation detectors just described. They also utilize faster, more forgiving photon detectors, covering different photon detection bandwidths. Radiator choices have broadened to include materials such as lithium fluoride, fused silica, and aerogel. Vacuum based photodetection systems (*e.g.*, single or multi anode PMTs, MCP PMTs, or hybrid photodiodes (HPD)) have become increasingly common (see Sec. 34.2 of this *Review*). They handle high rates, and can be used with a wide choice of radiators. Examples include (1) the SELEX RICH at Fermilab, which mirror focuses the Cherenkov photons from a neon radiator onto a camera array made of  $\sim 2000$  PMTs to separate hadrons over a wide momentum range (to well above 200 GeV/*c* for heavy hadrons); (2) the HERMES RICH at HERA, which mirror focuses photons from  $C_4F_{10}$  ( $n = 1.00137$ ) and aerogel ( $n = 1.0304$ ) radiators within the same volume onto a PMT camera array to separate hadrons in the momentum range from 2 to 15 GeV/*c*; and (3) the LHCb detector now running at the LHC. It uses two separate counters readout by hybrid PMTs. One volume, like HERMES, contains two radiators (aerogel and  $C_4F_{10}$ ) while the second volume contains  $CF_4$ . Photons are mirror focused onto detector arrays of HPDs to cover a  $\pi/K$  separation momentum range between 1 and 150 GeV/*c*. This device will be upgraded to deal with the higher luminosities provided by LHC after 2018 by modifying the optics and removing the aerogel radiator of the upstream RICH and replacing the Hybrid PMTs with multi-anode PMTs (MaPMTs).

Other fast detection systems that use solid cesium iodide (CsI) photocathodes or triethylamine (TEA) doping in proportional chambers are useful with certain radiator types and geometries. Examples include (1) the CLEO-III RICH at CESR that uses a LiF radiator with TEA doped proportional chambers; (2) the ALICE detector at the LHC that uses proximity focused liquid ( $C_6F_{14}$  radiators and solid CSI photocathodes (similar photodetectors have been used for several years by the HADES and COMPASS detectors), and the hadron blind detector (HBD) in the PHENIX detector at RHIC that couples a low index  $CF_4$  radiator to a photodetector based on electron multiplier (GEM) chambers with reflective CSI photocathodes [60].

A DIRC (Detection of Internally Reflected Cherenkov light) is a distinctive, compact RICH subtype first used in the BaBar detector [57,60]. A DIRC “inverts” the usual RICH principle for use of light from the radiator by collecting and imaging the total internally reflected light rather than the transmitted light. It utilizes the optical material of the radiator in two ways, simultaneously; first as a Cherenkov radiator, and second, as a light pipe. The magnitudes of the photon angles are preserved during transport by the flat, rectangular cross section radiators, allowing the photons to be efficiently transported to a detector outside the path of the particle where they may be imaged in up to three independent dimensions (the usual two in space and, due to the long photon paths lengths, one in time). Because the index of refraction in the radiator is large ( $\sim 1.48$  for fused silica), light collection efficiency is good, but the momentum range with good  $\pi/K$  separation is rather low. The BaBar DIRC range extends up to  $\sim 4$  GeV/*c*. It is plausible, but challenging, to extend it up to about 10 GeV/*c* with an improved design. New DIRC detectors are being developed that take advantage of the new, very fast, pixelated photodetectors becoming available, such as flat panel MaPMTs and MCP PMTs. They typically utilize either time imaging or mirror focused optics, or both, leading not only to a precision measurement of the Cherenkov angle, but in some cases, to a precise measurement of the particle TOF, and/or to correction of the chromatic dispersion in the radiator. Examples [60] include (1) the time of propagation (TOP) counter being fabricated for the BELLE-II upgrade at KEKB emphasizing precision timing for both Cherenkov imaging and TOF, which is scheduled for installation in 2016; (2) the full scale 3-dimensional imaging FDIRC prototype using the BaBar DIRC radiators which was designed for the SuperB detector at the Italian SuperB collider and uses precision timing not only for improving the angle reconstruction and TOF precision, but also to

correct the chromatic dispersion; (3) the DIRCs being developed for the PANDA detector at FAIR that use elegant focusing optics and fast timing; and (4) the TORCH proposal being developed for an LHCb upgrade after 2019 which uses DIRC imaging with fast photon detectors to provide particle separation via particle TOF over a path length of 9.5m.

### 34.6. Gaseous detectors

**34.6.1. Energy loss and charge transport in gases :** Revised March 2010 by F. Sauli (CERN) and M. Titov (CEA Saclay).

Gas-filled detectors localize the ionization produced by charged particles, generally after charge multiplication. The statistics of ionization processes having asymmetries in the ionization trails, affect the coordinate determination deduced from the measurement of drift time, or of the center of gravity of the collected charge. For thin gas layers, the width of the energy loss distribution can be larger than its average, requiring multiple sample or truncated mean analysis to achieve good particle identification. In the truncated mean method for calculating  $\langle dE/dx \rangle$ , the ionization measurements along the track length are broken into many samples and then a fixed fraction of high-side (and sometimes also low-side) values are rejected [64].

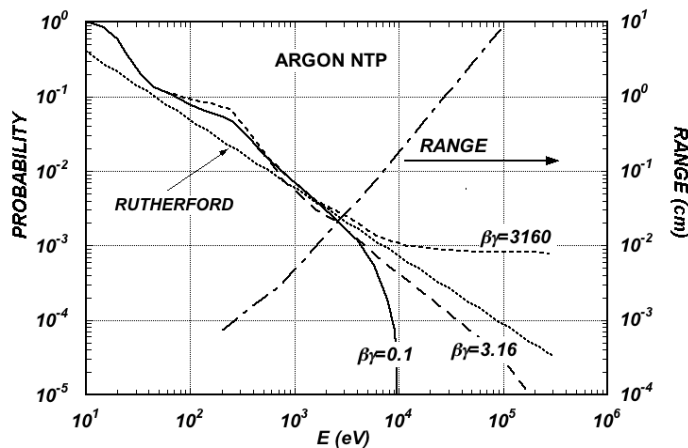
The energy loss of charged particles and photons in matter is discussed in Sec. 33. Table 34.5 provides values of relevant parameters in some commonly used gases at NTP (normal temperature, 20° C, and pressure, 1 atm) for unit-charge minimum-ionizing particles (MIPs) [65–71]. Values often differ, depending on the source, so those in the table should be taken only as approximate. For different conditions and for mixtures, and neglecting internal energy transfer processes (*e.g.*, Penning effect), one can scale the density,  $N_P$ , and  $N_T$  with temperature and pressure assuming a perfect gas law.

**Table 34.5:** Properties of noble and molecular gases at normal temperature and pressure (NTP: 20° C, one atm).  $E_X$ ,  $E_I$ : first excitation, ionization energy;  $W_I$ : average energy per ion pair;  $dE/dx|_{\min}$ ,  $N_P$ ,  $N_T$ : differential energy loss, primary and total number of electron-ion pairs per cm, for unit charge minimum ionizing particles.

Gas	Density, mg cm <sup>-3</sup>	$E_X$ eV	$E_I$ eV	$W_I$ eV	$dE/dx _{\min}$ keV cm <sup>-1</sup>	$N_P$ cm <sup>-1</sup>	$N_T$ cm <sup>-1</sup>
He	0.179	19.8	24.6	41.3	0.32	3.5	8
Ne	0.839	16.7	21.6	37	1.45	13	40
Ar	1.66	11.6	15.7	26	2.53	25	97
Xe	5.495	8.4	12.1	22	6.87	41	312
CH <sub>4</sub>	0.667	8.8	12.6	30	1.61	28	54
C <sub>2</sub> H <sub>6</sub>	1.26	8.2	11.5	26	2.91	48	112
iC <sub>4</sub> H <sub>10</sub>	2.49	6.5	10.6	26	5.67	90	220
CO <sub>2</sub>	1.84	7.0	13.8	34	3.35	35	100
CF <sub>4</sub>	3.78	10.0	16.0	54	6.38	63	120

When an ionizing particle passes through the gas it creates electron-ion pairs, but often the ejected electrons have sufficient energy to further ionize the medium. As shown in Table 34.5, the total number of electron-ion pairs ( $N_T$ ) is usually a few times larger than the number of primaries ( $N_P$ ).

The probability for a released electron to have an energy  $E$  or larger follows an approximate  $1/E^2$  dependence (Rutherford law), shown in Fig. 34.3 for Ar/CH<sub>4</sub> at NTP (dotted line, left scale). More detailed estimates taking into account the electronic structure of the medium are shown in the figure, for three values of the particle velocity factor  $\beta\gamma$  [66]. The dot-dashed line provides, on the right scale, the practical range of electrons (including scattering) of energy  $E$ . As an example, about 0.6% of released electrons have 1 keV or more energy, substantially increasing the ionization loss rate. The practical range of 1 keV electrons in argon (dot-dashed line, right scale) is 70  $\mu$ m and this can contribute to the error in the coordinate determination.



**Figure 34.3:** Probability of single collisions in which released electrons have an energy  $E$  or larger (left scale) and practical range of electrons in Ar/CH<sub>4</sub> (P10) at NTP (dot-dashed curve, right scale) [66].

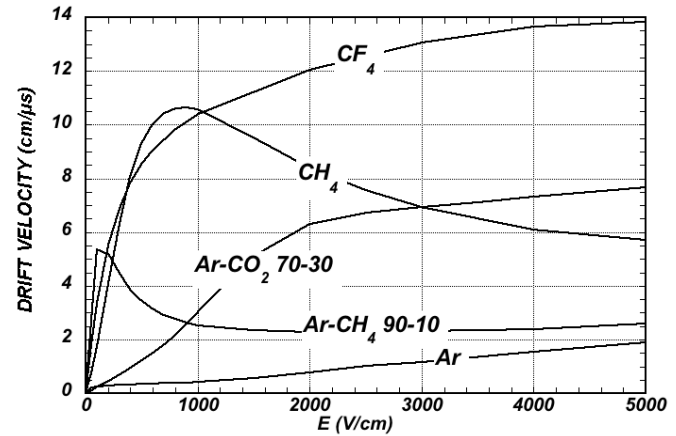
The number of electron-ion pairs per primary ionization, or cluster size, has an exponentially decreasing probability; for argon, there is about 1% probability for primary clusters to contain ten or more electron-ion pairs [67].

Once released in the gas, and under the influence of an applied electric field, electrons and ions drift in opposite directions and diffuse towards the electrodes. The scattering cross section is determined by the details of atomic and molecular structure. Therefore, the drift velocity and diffusion of electrons depend very strongly on the nature of the gas, specifically on the inelastic cross-section involving the rotational and vibrational levels of molecules. In noble gases, the inelastic cross section is zero below excitation and ionization thresholds. Large drift velocities are achieved by adding polyatomic gases (usually CH<sub>4</sub>, CO<sub>2</sub>, or CF<sub>4</sub>) having large inelastic cross sections at moderate energies, which results in “cooling” electrons into the energy range of the Ramsauer-Townsend minimum (at  $\sim 0.5$  eV) of the elastic cross-section of argon. The reduction in both the total electron scattering cross-section and the electron energy results in a large increase of electron drift velocity (for a compilation of electron-molecule cross sections see Ref. 68). Another principal role of the polyatomic gas is to absorb the ultraviolet photons emitted by the excited noble gas atoms. Extensive collections of experimental data [69] and theoretical calculations based on transport theory [70] permit estimates of drift and diffusion properties in pure gases and their mixtures. In a simple approximation, gas kinetic theory provides the drift velocity  $v$  as a function of the mean collision time  $\tau$  and the electric field  $E$ :  $v = eE\tau/m_e$  (Townsend’s expression). Values of drift velocity and diffusion for some commonly used gases at NTP are given in Fig. 34.4 and Fig. 34.5. These have been computed with the MAGBOLTZ program [71]. For different conditions, the horizontal axis must be scaled inversely with the gas density. Standard deviations for longitudinal ( $\sigma_L$ ) and transverse diffusion ( $\sigma_T$ ) are given for one cm of drift, and scale with the square root of the drift distance. Since the collection time is inversely proportional to the drift velocity, diffusion is less in gases such as CF<sub>4</sub> that have high drift velocities. In the presence of an external magnetic field, the Lorentz force acting on electrons between collisions deflects the drifting electrons and modifies the drift properties. The electron trajectories, velocities and diffusion parameters can be computed with MAGBOLTZ. A simple theory, the friction force model, provides an expression for the vector drift velocity  $\mathbf{v}$  as a function of electric and magnetic field vectors  $\mathbf{E}$  and  $\mathbf{B}$ , of the Larmor frequency  $\omega = eB/m_e$ , and of the mean collision time  $\tau$ :

$$\mathbf{v} = \frac{e}{m_e} \frac{\tau}{1 + \omega^2 \tau^2} \left( \mathbf{E} + \frac{\omega \tau}{B} (\mathbf{E} \times \mathbf{B}) + \frac{\omega^2 \tau^2}{B^2} (\mathbf{E} \cdot \mathbf{B}) \mathbf{B} \right) \quad (34.13)$$

To a good approximation, and for moderate fields, one can assume that the energy of the electrons is not affected by  $B$ , and use for  $\tau$

the values deduced from the drift velocity at  $B = 0$  (the Townsend expression). For  $\mathbf{E}$  perpendicular to  $\mathbf{B}$ , the drift angle to the relative to the electric field vector is  $\tan \theta_B = \omega \tau$  and  $v = (E/B)(\omega \tau / \sqrt{1 + \omega^2 \tau^2})$ . For parallel electric and magnetic fields, drift velocity and longitudinal diffusion are not affected, while the transverse diffusion can be strongly reduced:  $\sigma_T(B) = \sigma_T(B = 0) / \sqrt{1 + \omega^2 \tau^2}$ . The dotted line in Fig. 34.5 represents  $\sigma_T$  for the classic Ar/CH<sub>4</sub> (90:10) mixture at 4 T. Large values of  $\omega \tau \sim 20$  at 5 T are consistent with the measurement of diffusion coefficient in Ar/CF<sub>4</sub>/iC<sub>4</sub>H<sub>10</sub> (95:3:2). This reduction is exploited in time projection chambers (Sec. 34.6.5) to improve spatial resolution.



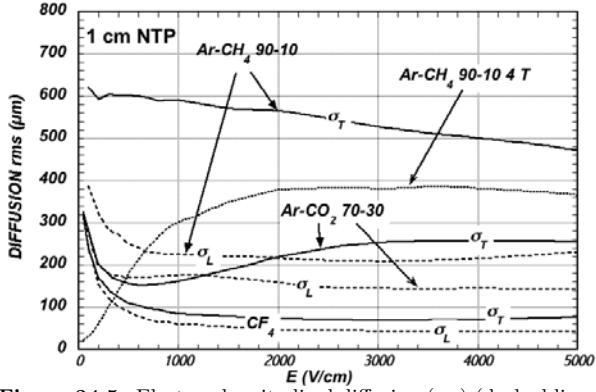
**Figure 34.4:** Computed electron drift velocity as a function of electric field in several gases at NTP and  $B = 0$  [71].

In mixtures containing electronegative molecules, such as O<sub>2</sub> or H<sub>2</sub>O, electrons can be captured to form negative ions. Capture cross-sections are strongly energy-dependent, and therefore the capture probability is a function of applied field. For example, the electron is attached to the oxygen molecule at energies below 1 eV. The three-body electron attachment coefficients may differ greatly for the same additive in different mixtures. As an example, at moderate fields (up to 1 kV/cm) the addition of 0.1% of oxygen to an Ar/CO<sub>2</sub> mixture results in an electron capture probability about twenty times larger than the same addition to Ar/CH<sub>4</sub>.

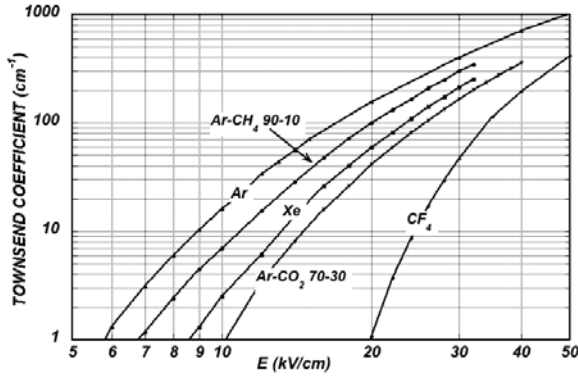
Carbon tetrafluoride is not electronegative at low and moderate fields, making its use attractive as drift gas due to its very low diffusion. However, CF<sub>4</sub> has a large electron capture cross section at fields above  $\sim 8$  kV/cm, before reaching avalanche field strengths. Depending on detector geometry, some signal reduction and resolution loss can be expected using this gas.

If the electric field is increased sufficiently, electrons gain enough energy between collisions to ionize molecules. Above a gas-dependent threshold, the mean free path for ionization,  $\lambda_i$ , decreases exponentially with the field; its inverse,  $\alpha = 1/\lambda_i$ , is the first Townsend coefficient. In wire chambers, most of the increase of avalanche particle density occurs very close to the anode wires, and a simple electrostatic consideration shows that the largest fraction of the detected signal is due to the motion of positive ions receding from the wires. The electron component, although very fast, contributes very little to the signal. This determines the characteristic shape of the detected signals in the proportional mode: a fast rise followed by a gradual increase. The slow component, the so-called “ion tail” that limits the time resolution of the detector, is usually removed by differentiation of the signal. In uniform fields,  $N_0$  initial electrons multiply over a length  $x$  forming an electron avalanche of size  $N = N_0 e^{\alpha x}$ ;  $N/N_0$  is the gain of the detector. Fig. 34.6 shows examples of Townsend coefficients for several gas mixtures, computed with MAGBOLTZ [71].

Positive ions released by the primary ionization or produced in the avalanches drift and diffuse under the influence of the electric field. Negative ions may also be produced by electron attachment to gas molecules. The drift velocity of ions in the fields encountered in



**Figure 34.5:** Electron longitudinal diffusion ( $\sigma_L$ ) (dashed lines) and transverse diffusion ( $\sigma_T$ ) (full lines) for 1 cm of drift at NTP and  $B = 0$ . The dotted line shows  $\sigma_T$  for the P10 mixture at 4 T [71].



**Figure 34.6:** Computed first Townsend coefficient  $\alpha$  as a function of electric field in several gases at NTP [71].

gaseous detectors (up to few kV/cm) is typically about three orders of magnitude less than for electrons. The ion mobility  $\mu$ , the ratio of drift velocity to electric field, is constant for a given ion type up to very high fields. Values of mobility at NTP for ions in their own and other gases are given in Table 34.6 [72]. For different temperatures and pressures, the mobility can be scaled inversely with the density assuming an ideal gas law. For mixtures, due to a very effective charge transfer mechanism, only ions with the lowest ionization potential survive after a short path in the gas. Both the lateral and transverse diffusion of ions are proportional to the square root of the drift time, with a coefficient that depends on temperature but not on the ion mass. Accumulation of ions in the gas drift volume may induce field distortions (see Sec. 34.6.5).

**Table 34.6:** Mobility of ions in gases at NTP [72].

Gas	Ion	Mobility $\mu$ ( $\text{cm}^2 \text{V}^{-1} \text{s}^{-1}$ )
He	$\text{He}^+$	10.4
Ne	$\text{Ne}^+$	4.7
Ar	$\text{Ar}^+$	1.54
Ar/ $\text{CH}_4$	$\text{CH}_4^+$	1.87
Ar/ $\text{CO}_2$	$\text{CO}_2^+$	1.72
$\text{CH}_4$	$\text{CH}_4^+$	2.26
$\text{CO}_2$	$\text{CO}_2^+$	1.09

**34.6.2. Multi-Wire Proportional and Drift Chambers :** Revised March 2010 by Fabio Sauli (CERN) and Maxim Titov (CEA Saclay).

Single-wire counters that detect the ionization produced in a gas by a charged particle, followed by charge multiplication and collection around a thin wire have been used for decades. Good energy resolution is obtained in the proportional amplification mode, while very large saturated pulses can be detected in the streamer and Geiger modes [3].

Multiwire proportional chambers (MWPCs) [73,74], introduced in the late '60's, detect, localize and measure energy deposit by charged particles over large areas. A mesh of parallel anode wires at a suitable potential, inserted between two cathodes, acts almost as a set of independent proportional counters (see Fig. 34.7a). Electrons released in the gas volume drift towards the anodes and produce avalanches in the increasing field. Analytic expressions for the electric field can be found in many textbooks. The fields close to the wires  $E(r)$ , in the drift region  $E_D$ , and the capacitance  $C$  per unit length of anode wire are approximately given by

$$E(r) = \frac{CV_0}{2\pi\epsilon_0} \frac{1}{r} \quad E_D = \frac{CV_0}{2\epsilon_0 s} \quad C = \frac{2\pi\epsilon_0}{\pi(\ell/s) - \ln(2\pi a/s)}, \quad (34.14)$$

where  $r$  is the distance from the center of the anode,  $s$  the wire spacing,  $\ell$  and  $V_0$  the distance and potential difference between anode and cathode, and  $a$  the anode wire radius.

Because of electrostatic forces, anode wires are in equilibrium only for a perfect geometry. Small deviations result in forces displacing the wires alternatively below and above the symmetry plane, sometimes with catastrophic results. These displacement forces are countered by the mechanical tension of the wire, up to a maximum unsupported stable length,  $L_M$  [64], above which the wire deforms:

$$L_M = \frac{s}{CV_0} \sqrt{4\pi\epsilon_0 T_M} \quad (34.15)$$

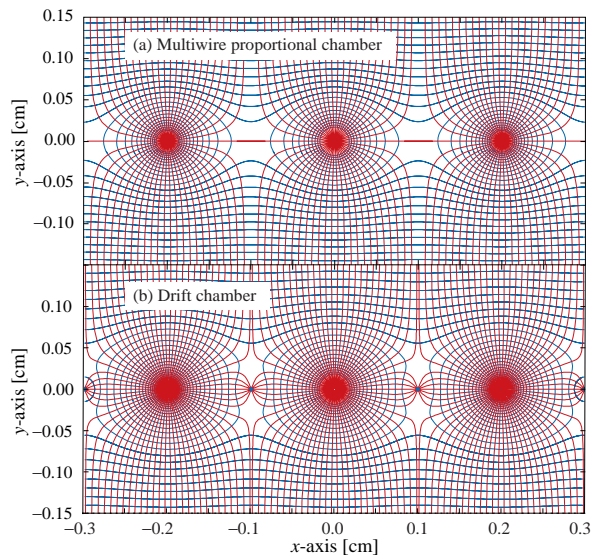
The maximum tension  $T_M$  depends on the wire diameter and modulus of elasticity. Table 34.7 gives approximate values for tungsten and the corresponding maximum stable wire length under reasonable assumptions for the operating voltage ( $V_0 = 5 \text{ kV}$ ) [75]. Internal supports and spacers can be used in the construction of longer detectors to overcome limits on the wire length imposed by Eq. (34.15).

**Table 34.7:** Maximum tension  $T_M$  and stable unsupported length  $L_M$  for tungsten wires with spacing  $s$ , operated at  $V_0 = 5 \text{ kV}$ . No safety factor is included.

Wire diameter ( $\mu\text{m}$ )	$T_M$ (newton)	$s$ (mm)	$L_M$ (cm)
10	0.16	1	25
20	0.65	2	85

Detection of charge on the wires over a predefined threshold provides the transverse coordinate to the wire with an accuracy comparable to that of the wire spacing. The coordinate along each wire can be obtained by measuring the ratio of collected charge at the two ends of resistive wires. Making use of the charge profile induced on segmented cathodes, the so-called center-of gravity (COG) method, permits localization of tracks to sub-mm accuracy. Due to the statistics of energy loss and asymmetric ionization clusters, the position accuracy is  $\sim 50 \mu\text{m}$  rms for tracks perpendicular to the wire plane, but degrades to  $\sim 250 \mu\text{m}$  at  $30^\circ$  to the normal [76]. The intrinsic bi-dimensional characteristic of the COG readout has found numerous applications in medical imaging.

Drift chambers, developed in the early '70's, can be used to estimate the longitudinal position of a track by exploiting the arrival time of electrons at the anodes if the time of interaction is known [77]. The distance between anode wires is usually several cm, allowing coverage of large areas at reduced cost. In the original design, a thicker wire



**Figure 34.7:** Electric field lines and equipotentials in (a) a multiwire proportional chamber and (b) a drift chamber.

(the field wire) at the proper voltage, placed between the anode wires, reduces the field at the mid-point between anodes and improves charge collection (Fig. 34.7b). In some drift chamber designs, and with the help of suitable voltages applied to field-shaping electrodes, the electric field structure is adjusted to improve the linearity of space-to-drift-time relation, resulting in better spatial resolution [78].

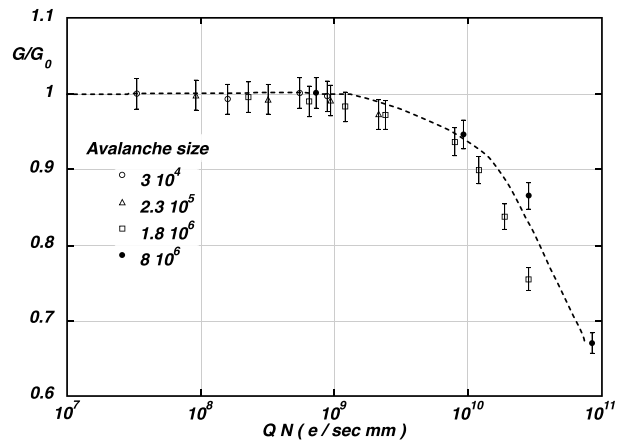
Drift chambers can reach a longitudinal spatial resolution from timing measurement of order  $100\ \mu\text{m}$  (rms) or better for minimum ionizing particles, depending on the geometry and operating conditions. However, a degradation of resolution is observed [79] due to primary ionization statistics for tracks close to the anode wires, caused by the spread in arrival time of the nearest ionization clusters. The effect can be reduced by operating the detector at higher pressures. Sampling the drift time on rows of anodes led to the concept of multiple arrays such as the multi-drift module [80] and the JET chamber [81]. A measurement of drift time, together with the recording of charge sharing from the two ends of the anode wires provides the coordinates of segments of tracks. The total charge gives information on the differential energy loss and is exploited for particle identification. The time projection chamber (TPC) [82] combines a measurement of drift time and charge induction on cathodes, to obtain excellent tracking for high multiplicity topologies occurring at moderate rates (see Sec. 34.6.5). In all cases, a good knowledge of electron drift velocity and diffusion properties is required. This has to be combined with the knowledge of the electric fields in the structures, computed with commercial or custom-developed software [71,83]. For an overview of detectors exploiting the drift time for coordinate measurement see Refs. 6 and 64.

Multiwire and drift chambers have been operated with a variety of gas fillings and operating modes, depending on experimental requirements. The so-called “Magic Gas,” a mixture of argon, isobutane and Freon [74], permits very high and saturated gains ( $\sim 10^6$ ). This gas mixture was used in early wire chambers, but was found to be susceptible to severe aging processes. With present-day electronics, proportional gains around  $10^4$  are sufficient for detection of minimum ionizing particles, and noble gases with moderate amounts of polyatomic gases, such as methane or carbon dioxide, are used.

Although very powerful in terms of performance, multi-wire structures have reliability problems when used in harsh or hard-to-access environments, since a single broken wire can disable the entire detector. Introduced in the '80's, straw and drift tube systems make use of large arrays of wire counters encased in individual enclosures, each acting as an independent wire counter [84]. Techniques for low-cost mass production of these detectors have been developed for large experiments, such as the Transition Radiation Tracker and the Drift Tubes arrays for CERN's LHC experiments [85].

### 34.6.3. High Rate Effects : Revised March 2010 by Fabio Sauli (CERN) and Maxim Titov (CEA Saclay).

The production of positive ions in the avalanches and their slow drift before neutralization result in a rate-dependent accumulation of positive charge in the detector. This may result in significant field distortion, gain reduction and degradation of spatial resolution. As shown in Fig. 34.8 [86], the proportional gain drops above a charge production rate around  $10^9$  electrons per second and mm of wire, independently of the avalanche size. For a proportional gain of  $10^4$  and 100 electrons per track, this corresponds to a particle flux of  $10^3\text{ s}^{-1}\text{ mm}^{-1}$  (1 kHz/mm<sup>2</sup> for 1 mm wire spacing).



**Figure 34.8:** Charge rate dependence of normalized gas gain  $G/G_0$  (relative to zero counting rate) in proportional thin-wire detectors [86].  $Q$  is the total charge in single avalanche;  $N$  is the particle rate per wire length.

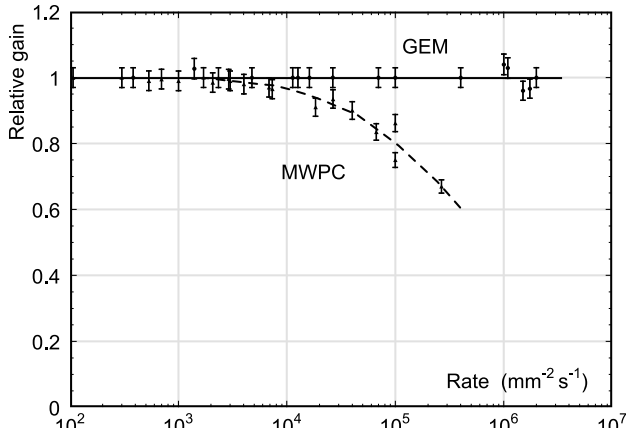
At high radiation fluxes, a fast degradation of detectors due to the formation of polymers deposits (aging) is often observed. The process has been extensively investigated, often with conflicting results. Several causes have been identified, including organic pollutants and silicone oils. Addition of small amounts of water in many (but not all) cases has been shown to extend the lifetime of the detectors. Addition of fluorinated gases (e.g.,  $\text{CF}_4$ ) or oxygen may result in an etching action that can overcome polymer formation, or even eliminate already existing deposits. However, the issue of long-term survival of gas detectors with these gases is controversial [87]. Under optimum operating conditions, a total collected charge of a few coulombs per cm of wire can usually be reached before noticeable degradation occurs. This corresponds, for one mm spacing and at a gain of  $10^4$ , to a total particle flux of  $\sim 10^{14}$  MIPs/cm<sup>2</sup>.

### 34.6.4. Micro-Pattern Gas Detectors : Revised March 2010 by Fabio Sauli (CERN) and Maxim Titov (CEA Saclay)

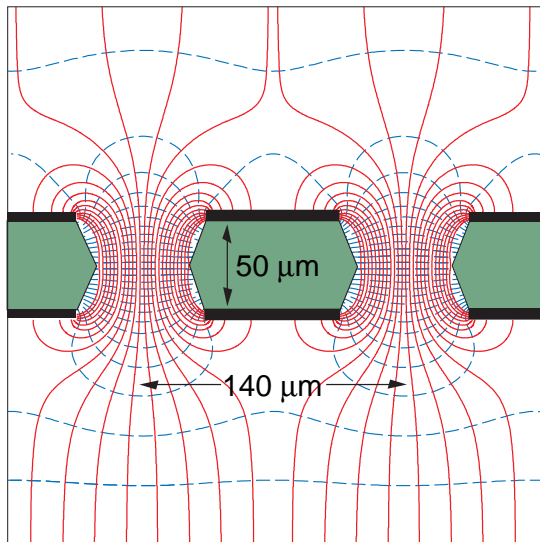
Despite various improvements, position-sensitive detectors based on wire structures are limited by basic diffusion processes and space charge effects to localization accuracies of  $50\text{--}100\ \mu\text{m}$  [88]. Modern photolithographic technology led to the development of novel Micro-Pattern Gas Detector (MPGD) concepts [89], revolutionizing cell size limitations for many gas detector applications. By using pitch size of a few hundred  $\mu\text{m}$ , an order of magnitude improvement in granularity over wire chambers, these detectors offer intrinsic high rate capability ( $> 10^6\text{ Hz/mm}^2$ ), excellent spatial resolution ( $\sim 30\ \mu\text{m}$ ), multi-particle resolution ( $\sim 500\ \mu\text{m}$ ), and single photo-electron time resolution in the ns range.

The Micro-Strip Gas Chamber (MSGC), invented in 1988, was the first of the micro-structure gas chambers [90]. It consists of a set of tiny parallel metal strips laid on a thin resistive support, alternatively connected as anodes and cathodes. Owing to the small anode-to-cathode distance ( $\sim 100\ \mu\text{m}$ ), the fast collection of positive ions reduces space charge build-up, and provides a greatly increased rate capability. Unfortunately, the fragile electrode structure of the MSGC turned out to be easily destroyed by discharges induced by

heavily ionizing particles [91]. Nevertheless, detailed studies of their properties, and in particular, on the radiation-induced processes leading to discharge breakdown, led to the development of the more powerful devices: GEM and Micromegas. These have improved reliability and radiation hardness. The absence of space-charge effects in GEM detectors at the highest rates reached so far and the fine granularity of MPGDs improve the maximum rate capability by more than two orders of magnitude (Fig. 34.9) [78,92]. Even larger rate capability has been reported for Micromegas [93].

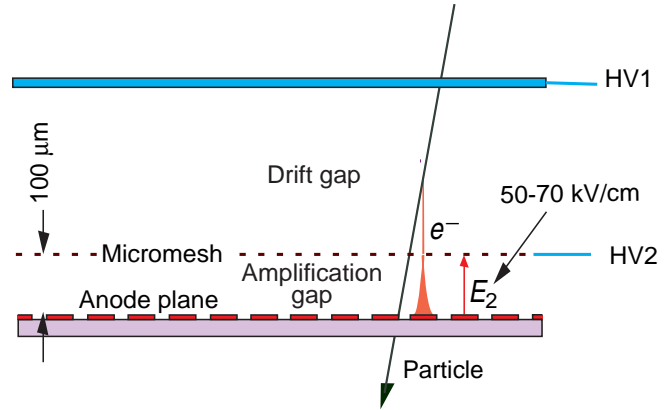


**Figure 34.9:** Normalized gas gain as a function of particle rate for MWPC [78] and GEM [92].



**Figure 34.10:** Schematic view and typical dimensions of the hole structure in the GEM amplification cell. Electric field lines (solid) and equipotentials (dashed) are shown.

The Gas Electron Multiplier (GEM) detector consists of a thin-foil copper-insulator-copper sandwich chemically perforated to obtain a high density of holes in which avalanches occur [94]. The hole diameter is typically between 25 μm and 150 μm, while the corresponding distance between holes varies between 50 μm and 200 μm. The central insulator is usually (in the original design) the polymer Kapton, with a thickness of 50 μm. Application of a potential difference between the two sides of the GEM generates the electric fields indicated in Fig. 34.10. Each hole acts as an independent proportional counter. Electrons released by the primary ionization particle in the upper conversion region (above the GEM foil) drift into the holes, where charge multiplication occurs in the high electric field (50–70 kV/cm). Most of avalanche electrons are transferred



**Figure 34.11:** Schematic drawing of the Micromegas detector.

into the gap below the GEM. Several GEM foils can be cascaded, allowing the multi-layer GEM detectors to operate at overall gas gain above 10<sup>4</sup> in the presence of highly ionizing particles, while strongly reducing the risk of discharges. This is a major advantage of the GEM technology [95]. Localization can then be performed by collecting the charge on a patterned one- or two-dimensional readout board of arbitrary pattern, placed below the last GEM.

The micro-mesh gaseous structure (Micromegas) is a thin parallel-plate avalanche counter, as shown in Fig. 34.11 [96]. It consists of a drift region and a narrow multiplication gap (25–150 μm) between a thin metal grid (micromesh) and the readout electrode (strips or pads of conductor printed on an insulator board). Electrons from the primary ionization drift through the holes of the mesh into the narrow multiplication gap, where they are amplified. The electric field is homogeneous both in the drift (electric field ~1 kV/cm) and amplification (50–70 kV/cm) gaps. In the narrow multiplication region, gain variations due to small variations of the amplification gap are approximately compensated by an inverse variation of the amplification coefficient, resulting in a more uniform gain. The small amplification gap produces a narrow avalanche, giving rise to excellent spatial resolution: 12 μm accuracy, limited by the micro-mesh pitch, has been achieved for MIPs, as well as very good time resolution and energy resolution (~12% FWHM with 6 keV x rays) [97].

The performance and robustness of GEM and Micromegas have encouraged their use in high-energy and nuclear physics, UV and visible photon detection, astroparticle and neutrino physics, neutron detection and medical physics. Most structures were originally optimized for high-rate particle tracking in nuclear and high-energy physics experiments. COMPASS, a high-luminosity experiment at CERN, pioneered the use of large-area (~40 × 40 cm<sup>2</sup>) GEM and Micromegas detectors close to the beam line with particle rates of 25 kHz/mm<sup>2</sup>. Both technologies achieved a tracking efficiency of close to 100% at gas gains of about 10<sup>4</sup>, a spatial resolution of 70–100 μm and a time resolution of ~10 ns. GEM detectors are also used for triggering in the LHCb Muon System and for tracking in the TOTEM Telescopes. Both GEM and Micromegas devices are foreseen for the upgrade of the LHC experiments and for one of the readout options for the Time Projection Chamber (TPC) at the International Linear Collider (ILC). The development of new fabrication techniques—“bulk” Micromegas technology [98] and single-mask GEMs [99]—is a big step toward industrial production of large-size MPGDs. In some applications requiring very large-area coverage with moderate spatial resolution, coarse macro-patterned detectors, such as Thick GEMs (THGEM) [100] or patterned resistive-plate devices [101] might offer economically interesting solutions.

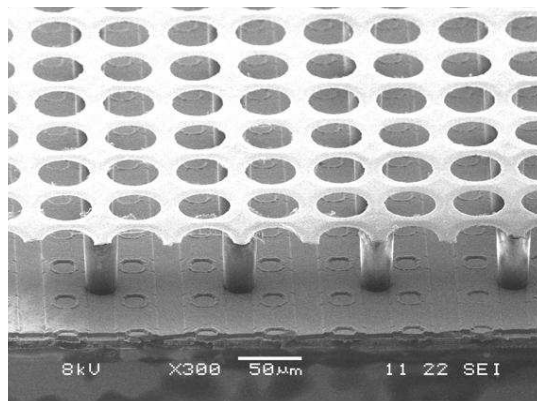
Sensitive and low-noise electronics enlarge the range of the MPGD applications. Recently, the GEM and Micromegas detectors were read out by high-granularity (~50 μm pitch) CMOS chips assembled directly below the GEM or Micromegas amplification structures [102]. These detectors use the bump-bonding pads of a pixel chip as an integrated charge collecting anode. With this arrangement signals are



induced at the input gate of a charge-sensitive preamplifier (top metal layer of the CMOS chip). Every pixel is then directly connected to the amplification and digitization circuits, integrated in the underlying active layers of the CMOS technology, yielding timing and charge measurements as well as precise spatial information in 3D.

The operation of a MPGD with a Timepix CMOS chip has demonstrated the possibility of reconstructing 3D-space points of individual primary electron clusters with  $\sim 30\ \mu\text{m}$  spatial resolution and event-time resolution with nanosecond precision. This has become indispensable for tracking and triggering and also for discriminating between ionizing tracks and photon conversions. The GEM, in conjunction with a CMOS ASIC,\* can directly view the absorption process of a few keV x-ray quanta and simultaneously reconstruct the direction of emission, which is sensitive to the x-ray polarization. Thanks to these developments, a micro-pattern device with finely segmented CMOS readout can serve as a high-precision “electronic bubble chamber.” This may open new opportunities for x-ray polarimeters, detection of weakly interacting massive particles (WIMPs) and axions, Compton telescopes, and 3D imaging of nuclear recoils.

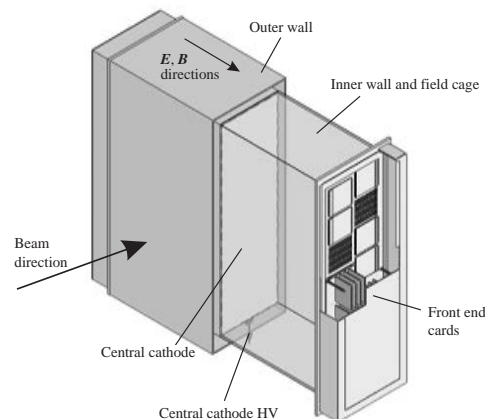
An elegant solution for the construction of the Micromegas with pixel readout is the integration of the amplification grid and CMOS chip by means of an advanced “wafer post-processing” technology [103]. This novel concept is called “Ingrid” (see Fig. 34.12). With this technique, the structure of a thin ( $1\ \mu\text{m}$ ) aluminum grid is fabricated on top of an array of insulating pillars, which stands  $\sim 50\ \mu\text{m}$  above the CMOS chip. The sub- $\mu\text{m}$  precision of the grid dimensions and avalanche gap size results in a uniform gas gain. The grid hole size, pitch and pattern can be easily adapted to match the geometry of any pixel readout chip.



**Figure 34.12:** Photo of the Micromegas “Ingrid” detector. The grid holes can be accurately aligned with readout pixels of CMOS chip. The insulating pillars are centered between the grid holes, thus avoiding dead regions.

Recent developments in radiation hardness research with state-of-the-art MPGDs are reviewed in Ref. 104. Earlier aging studies of GEM and Micromegas concepts revealed that they might be even less vulnerable to radiation-induced performance degradation than standard silicon microstrip detectors.

The RD51 collaboration was established in 2008 to further advance technological developments of micro-pattern detectors and associated electronic-readout systems for applications in basic and applied research [105].



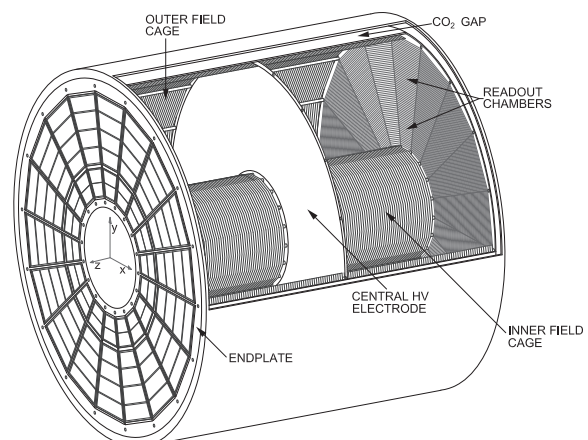
**Figure 34.14:** One of the 3 TPC modules for the near detector of the T2K experiment [107]. The size is  $2 \times 2 \times 0.8\ \text{m}^3$ . Micromegas devices are used for gas amplification and readout.

**34.6.5. Time-projection chambers :** Written August 2015 by C. Lippmann (GSI Helmholtzzentrum für Schwerionenforschung, Darmstadt, Germany)

The Time Projection Chamber (TPC) concept was invented by David Nygren in the late 1970’s [82]. It consists of a cylindrical or square field cage filled with a detection medium that is usually a gas or a liquid. Charged particles produce tracks of ionization electrons that drift in a uniform electric field towards a position-sensitive amplification stage which provides a 2D projection of the particle trajectories. The third coordinate can be calculated from the arrival times of the drifted electrons. The start for this drift time measurement is usually derived from an external detector, e.g. a fast interaction trigger detector.

This section focuses on the gas-filled TPCs that are typically used in particle or nuclear physics experiments at accelerators due to their low material budget. For neutrino physics (Sec. 34.10) or for detecting rare events (Sec. 35.4), on the contrary, usually high density and large active mass are required, and a liquid detection medium is favored.

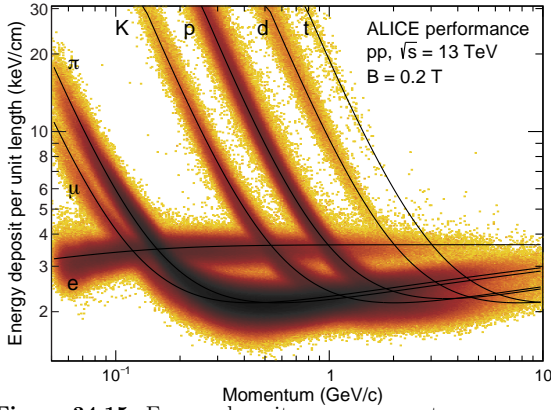
The TPC enables full 3D measurements of charged particle tracks, which gives it a distinct advantage over other tracking detector designs which record information only in two-dimensional detector planes and have less overall segmentation. This advantage is often exploited for pattern recognition in events with large numbers of particles, e.g. heavy-ion collisions. Two examples of modern large-volume gaseous TPCs are shown in Fig. 34.13 and Fig. 34.14.



**Figure 34.13:** Schematic view of the ALICE TPC [106]. The drift volume with 5 m diameter is divided into two halves, each providing 2.5 m drift length.

\* Application Specific Integrated Circuit

Identification of the charged particles crossing the TPC is possible by simultaneously measuring their momentum and specific energy deposit through ionisation ( $dE/dx$ ). The momentum, as well as the charge sign, are calculated from a helix fit to the particle trajectory in the presence of a magnetic field (typically parallel to the drift field). For this application, precise spatial measurements in the plane transverse to the magnetic field are most important. The specific energy deposit is estimated from many charge measurements along the particle trajectory (e.g. one measurement per anode wire or per row of readout pads). As the charge collected per readout segment depends on the track angle and on the ambient conditions, the measured values are corrected for the effective length of the track segments and for variations of the gas temperature and pressure. The most probable value of the corrected signal amplitudes provides the best estimator for the specific energy deposit (see Sec. 33.2.3); it is usually approximated by the truncated mean, i.e. the average of the 50%-70% smallest values. The resulting particle identification performance is illustrated in Fig. 34.15, for the ALICE TPC.



**Figure 34.15:** Energy deposit versus momentum measured in the ALICE TPC [108].

The dependence of the achievable energy resolution on the number of measurements  $N$ , on the thickness of the sampling layers  $t$ , and on the gas pressure  $P$  can be estimated using an empirical formula [109]:

$$\sigma_{dE/dx} = 0.41 N^{-0.43} (tP)^{-0.32}. \quad (34.16)$$

Typical values at nominal pressure are  $\sigma_{dE/dx} = 4.5$  to  $7.5\%$ , with  $t = 0.4$  to  $1.5$  cm and  $N = 40$  up to more than 300. Due to the high gas pressure of 8.5 bar, the resolution achieved with the PEP-4/9 TPC was an unprecedented 3% [110].

The greatest challenges for a large TPC are due to the length of the drift of up to several meters. In particular, it can make the device sensitive to small distortions in the electric field. Such distortions can arise from a number of sources, e.g. imperfections in the field cage construction or the presence of ions in the drift volume. The electron drift in a TPC in the presence of a magnetic field is defined by Eq. (34.13). The  $\mathbf{E} \times \mathbf{B}$  term of Eq. (34.13) vanishes for perfectly aligned electric and magnetic fields, which can however be difficult to achieve in practice. Furthermore, the electron drift depends on the  $\omega\tau$  factor, which is defined by the chosen gas mixture and magnetic field strength. The electrons will tend to follow the magnetic field lines for  $\omega\tau > 1$  or the electric field lines for  $\omega\tau < 1$ . The former mode of operation makes the TPC less sensitive to non-uniformities of the electric field, which is usually desirable.

The drift of the ionization electrons is superposed with a random diffusion motion which degrades their position information. The ultimate resolution of a single position measurement is limited to around

$$\sigma_x = \frac{\sigma_D \sqrt{L}}{\sqrt{n}}, \quad (34.17)$$

where  $\sigma_D$  is the transverse diffusion coefficient for 1 cm drift,  $L$  is the drift length in cm and  $n$  is the effective number of electrons collected.

Without a magnetic field,  $\sigma_{D,B=0} \sqrt{L}$  is typically a few mm after a drift of  $L = 100$  cm. However, in a strong magnetic field parallel to the drift field, a large value of  $\omega\tau$  can significantly reduce diffusion:

$$\frac{\sigma_{D,B>0}}{\sigma_{D,B=0}} = \frac{1}{\sqrt{1 + \omega^2 \tau^2}}. \quad (34.18)$$

This factor can reach values of up to 10. In practice, the final resolution limit due to diffusion will typically be around  $\sigma_x = 100 \mu\text{m}$ .

The drift and diffusion of electrons depend strongly on the nature of the gas that is used. The optimal gas mixture varies according to the environment in which the TPC will operate. In all cases, the oxygen concentration must be kept very low (few ten parts per million in a large TPC) in order to avoid electron loss through attachment. Ideally, the drift velocity should depend only weakly on the electric field at the nominal operating condition. The classic Ar/CH<sub>4</sub> (90:10) mixture, known as P10, has a drift velocity maximum of 5 cm/ $\mu\text{s}$  at an electric field of only 125 V/cm (Fig. 34.4). In this regime, the electron arrival time is not affected by small variations in the ambient conditions. Moreover, low electric fields simplify the design and operation of the field cage. The mixture has a large transverse diffusion at  $B = 0$ , but this can be reduced significantly in a strong magnetic field due to the relatively large value of  $\omega\tau$ .

For certain applications, organic gases like CH<sub>4</sub> are not desirable, since they may cause aging. An alternative is to replace CH<sub>4</sub> with CO<sub>2</sub>. An Ar/CO<sub>2</sub> (90:10) mixture features a low transverse diffusion at all magnetic field strengths, but does not provide a saturated drift velocity for the typical electric fields used in TPCs (up to a few 100 V/cm), so it is quite sensitive to the ambient conditions. Freon admixtures like CF<sub>4</sub> can be an attractive option for a TPC as well, since the resulting gas mixtures provide high drift velocities at low electric fields. However, the use of CF<sub>4</sub> always needs to be thoroughly validated for compatibility with all materials of the detector and the gas system.

Historically, the amplification stages used in gaseous TPCs have been planes of anode wires operated in proportional mode. The performance is limited by effects related to the feature size of a few mm (wire spacing). Since near the wires the electric and magnetic fields are not parallel, the incoming ionisation electrons are displaced in the direction of the wires (“wire  $\mathbf{E} \times \mathbf{B}$  effect”), which degrades the resolution. The smaller feature sizes of Micro-Pattern Gas Detectors (MPGDs) like GEMs and Micromegas lead to many advantages as compared to wire planes (see Sec. 34.6.4). In particular,  $\mathbf{E} \times \mathbf{B}$  effects in the amplification stage are much smaller. Moreover, the signal induction process in MPGDs leads to a very narrow pad response, allowing for a much finer segmentation and improving the separation of two nearby tracks. Combinations of MPGDs with silicon sensors have resulted in the highest granularity readout systems so far (see Sec. 34.6.4). These devices make it possible to count the number of ionization clusters along the length of a track, which can, in principle, improve the particle identification capability. However, the big challenge for such a system is the huge number of read-out channels for a TPC of a typical size.

The accumulation of the positive ions created by the ionization from the particle tracks can lead to time-dependent distortions of the drift field. Due to their small drift velocity, ions from many events may coexist in the drift volume. To reduce the effect of such a build-up of space charge, Argon can be replaced by Neon as the main component of the gas mixture. Neon features a lower number of ionisation electrons per unit of track length (see Table 34.5) and a higher ion mobility (see Table 34.6).

Of much greater concern are the ions produced in the gas amplification stage. In order to prevent them from entering the drift volume, large TPCs built until now usually have a gating grid. The gating grid can be switched to transparent mode (usually in the presence of an interaction trigger) to allow the ionization electrons to pass into the amplification region. After all electrons have reached the amplification region, it is usually closed such that it is rendered opaque to electrons and ions.

Alternatively, new readout schemes are being developed using MPGDs. These can be optimized in a way that they release many

fewer positive ions than wire planes operating at the same effective gain. This is an exciting possibility for future TPCs.

**34.6.6. Transition radiation detectors (TRD's) :** Revised August 2013 by P. Nevski (BNL) and A. Romaniouk (Moscow Eng. & Phys. Inst.)

Transition radiation (TR) x-rays are produced when a highly relativistic particle ( $\gamma \gtrsim 10^3$ ) crosses a refractive index interface, as discussed in Sec. 33.7. The x-rays, ranging from a few keV to a few dozen keV or more, are emitted at a characteristic angle  $1/\gamma$  from the particle trajectory. Since the TR yield is about 1% per boundary crossing, radiation from multiple surface crossings is used in practical detectors. In the simplest concept, a detector module might consist of low- $Z$  foils followed by a high- $Z$  active layer made of proportional counters filled with a Xe-rich gas mixture. The atomic number considerations follow from the dominant photoelectric absorption cross section per atom going roughly as  $Z^n/E_x^3$ , where  $n$  varies between 4 and 5 over the region of interest, and the x-ray energy is  $E_x$ .<sup>\*</sup> To minimize self-absorption, materials such as polypropylene, Mylar, carbon, and (rarely) lithium are used as radiators. The TR signal in the active regions is in most cases superimposed upon the particle ionization losses, which are proportional to  $Z$ .

The TR intensity for a single boundary crossing always increases with  $\gamma$ , but, for multiple boundary crossings, interference leads to saturation above a Lorentz factor  $\gamma_{\text{sat}} = 0.6 \omega_1 \sqrt{\ell_1 \ell_2} / c$  [111], where  $\omega_1$  is the radiator material plasma frequency,  $\ell_1$  is its thickness, and  $\ell_2$  the spacing. In most of the detectors used in particle physics the radiator parameters are chosen to provide  $\gamma_{\text{sat}} \approx 2000$ . Those detectors normally work as threshold devices, ensuring the best electron/pion separation in the momentum range  $1 \text{ GeV}/c \lesssim p \lesssim 150 \text{ GeV}/c$ .

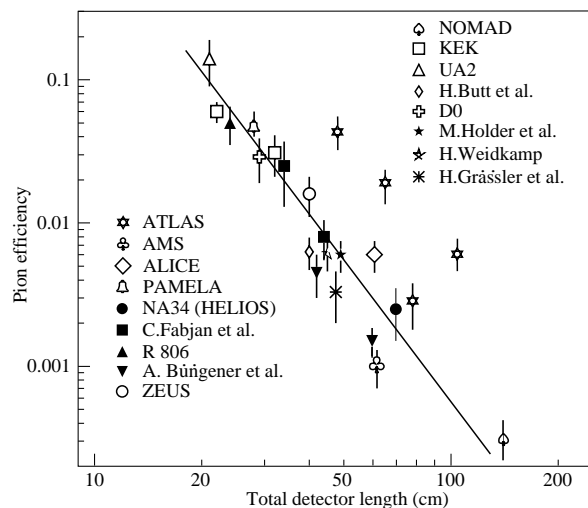
One can distinguish two design concepts—“thick” and “thin” detectors:

1. The radiator, optimized for a minimum total radiation length at maximum TR yield and total TR absorption, consists of few hundred foils (for instance 300 20  $\mu\text{m}$  thick polypropylene foils). Most of the TR photons are absorbed in the radiator itself. To maximise the number of TR photons reaching the detector, part of the radiator far from the active layers is often made of thicker foils, which shifts the x-ray spectrum to higher energies. The detector thickness, about 2-4 cm for Xe-filled gas chambers, is optimized to absorb the incoming x-ray spectrum. A classical detector is composed of several similar modules which respond nearly independently. Such detectors were used in the UA2, NA34 and other experiments [112], and are being used in the ALICE experiment [113], [114].
2. In other TRD concepts a fine granular radiator/detector structure exploits the soft part of the TR spectrum more efficiently and thereby may act also as an integral part of the tracking detector. This can be achieved, for instance, by distributing small-diameter straw-tube detectors uniformly or in thin layers throughout the radiator material (foils or fibers). Even with a relatively thin radiator stack, radiation below 5 keV is mostly lost in the radiators themselves. However for photon energies above this value, the absorption is reduced and the radiation can be registered by several consecutive detector layers, thus creating a strong TR build-up effect. This approach allows to realise TRD as an integral part of the tracking detector. Descriptions of detectors using this approach can be found in both accelerator and space experiments [113,114]. For example, in the ATLAS TR tracker (TRT), charged particles cross about 35 effective straw tube layers embedded in the radiator material [113]. The effective thickness of the Xe gas per straw is about 2.2 mm and the average number of foils per straw is about 40 with an effective foil thickness of about 18  $\mu\text{m}$ .

Both TR photon absorption and the TR build-up significantly affect the detector performance. Although the values mentioned above are

<sup>\*</sup> Photon absorption coefficients for the elements (via a NIST link), and  $dE/dx_{\text{min}}$  and plasma energies for many materials are given in [pdg.lbl.gov/AtomicNuclearProperties](http://pdg.lbl.gov/AtomicNuclearProperties).

typical for most of the plastic radiators used with Xe-based detectors, they vary significantly depending on the detector parameters: radiator material, thickness and spacing, the geometry and position of the sensitive chambers, etc. Thus careful simulations are usually needed to build a detector optimized for a particular application. For TRD simulation stand-alone codes based on GEANT3 program were usually used (P.Nevski in [113]). TR simulation is now available in GEANT4 [118]. The most recent version of it (starting from release 9.5) shows a reasonable agreement with data (S. Furletov in [114] and [115]).



**Figure 34.16:** Pion efficiency measured (or predicted) for different TRDs as a function of the detector length for a fixed electron efficiency of 90%. The plot is taken from [112]. Results from more recent detectors are added from [113] and [114].

The discrimination between electrons and pions can be based on the charge deposition measured in each detection module, on the number of clusters – energy depositions observed above an optimal threshold (usually it is 5–7 keV), or on more sophisticated methods such as analyzing the pulse shape as a function of time. The total energy measurement technique is more suitable for thick gas volumes, which absorb most of the TR radiation and where the ionization loss fluctuations are small. The cluster-counting method works better for detectors with thin gas layers, where the fluctuations of the ionization losses are big. Cluster-counting replaces the Landau-Vavilov distribution of background ionization energy losses with the Poisson statistics of  $\delta$ -electrons, responsible for the distribution tails. The latter distribution is narrower than the Landau-Vavilov distribution. In practice, most of the experiments use a likelihood method, which exploits detailed knowledge of the detector response for different particles and gives the best separation. The more parameters that are considered, the better separation power. The recent results of the TRD in the AMS experiment is a good example. In the real experiment the rejection power is better by almost one order of magnitude than that obtained in the beam test if stringent criteria for track selection are applied (see T. Kirn *et al.* in [114]). Another example is a neural network method used by the ALICE TRD (ALICE point in Fig. 34.16) which gives another factor of 2–3 in rejection power with respect to the likelihood method [116]).

The major factor in the performance of any TRD is its overall length. This is illustrated in Fig. 34.16, which shows, for a variety of detectors, the pion efficiency at a fixed electron efficiency of 90% as a function of the overall detector length. TRD performance depends on particle energy and in this figure the experimental data, covering a range of particle energies from 1 GeV to 40 GeV, are rescaled to an energy of 10 GeV when possible. Phenomenologically, the rejection power against pions increases as  $5 \cdot 10^{L/38}$ , where the range of validity is  $L \approx 20\text{--}100 \text{ cm}$ . Apart from the beam energy variations, the observed scattering of the points in the plot reflects how effectively the detector space is used and how well the exact response to different



particles is taken into account in the analysis. For instance, the ATLAS TRT was built as a compromise between TR and tracking requirements; that is why the test-beam prototype result (lower point) is better than the real TRT performance at the LHC shown in Fig. 34.16 for different regions in the detector (in agreement with MC).

In most cases, recent TRDs combine particle identification with charged-track measurement in the same detector [113,114,117]. This is particularly important for collider experiments, where the available space for the inner detector is very limited. For a modest increase of the radiation length due to the radiator ( $\sim 4\%$   $X_0$ ), a significant enhancement of the electron identification was obtained in the case of the ATLAS TRT. The combination of the two detector functions provides a powerful tool for electron identification even at very high particle densities.

In addition to the enhancement of the electron identification, one of the most important roles of the TRDs in the collider experiments is their participation in different trigger and data analysis algorithms. The ALICE experiment [114] is a good example of the use of the TRD in a First Level Trigger. In the ATLAS experiment, the TRT information is used in the High Level Trigger (HLT) algorithms. With continuous increase of instantaneous luminosity, the electron trigger output rate becomes so high, that a significant increase of the calorimeter energy threshold is required to keep it at an acceptable level. For luminosities above  $2 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$  at the LHC this will affect the trigger efficiency of very important physics channels (e.g.  $W \rightarrow e\nu$  inclusive decay). Even a very soft TR cut at HLT level, which preserves high electron efficiency (98%), allows to maintain a high trigger efficiency and its purity for physics events with a single electron in a final state. TRT also plays a crucial role in the studies where an electron suppression is required (e.g. hadronic mode of  $\tau$ -decays). TR information is a completely independent tool for electron identification and allows to study systematic uncertainties of other electron reconstruction methods.

Electron identification is not the only TRD application. Recent TRDs for particle astrophysics are designed to directly measure the Lorentz factor of high-energy nuclei by using the quadratic dependence of the TR yield on nuclear charge; see Cherry and Müller papers in [113]. The radiator configuration ( $\ell_1, \ell_2$ ) is tuned to extend the TR yield rise up to  $\gamma \lesssim 10^5$  using the more energetic part of the TR spectrum (up to 100 keV). Large density radiator materials (such as Al) are the best for this purpose. Direct absorption of the TR-photons of these energies with thin detectors becomes problematic and TR detection methods based on Compton scattering have been proposed to use (M. Cherry in [113], [114]).

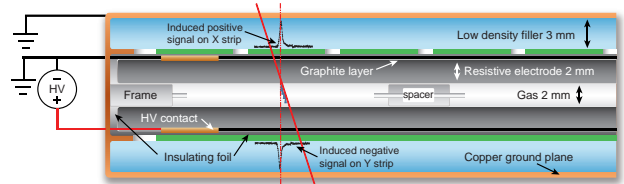
In all cases to-date, the radiator properties have been the main limiting factor for the TRDs, and for future progress in this field, it is highly important to develop effective and compact radiators. By now, all traditional materials have been studied extensively, so new technologies must be invented. The properties of all radiators are defined by one basic parameter which is the plasma frequency of the radiator material –  $\omega_1 \sim 1/m_e$  (see Eq. (33.48)). In semiconductor materials, a quantum mechanical treatment of the electron binding to the lattice leads to a small effective electron mass and correspondingly to large values of  $\omega_1$ . All semiconductor materials have large  $Z$  and may not be good candidates as TR radiators, but new materials, such as graphene, may offer similar features at much lower  $Z$  (M. Cherry in [114]). It might even be possible to produce graphene-based radiators with the required  $\omega_1$  value. One should take into account that TR cutoff energy –  $E_c \sim \omega_1 \gamma$  and 95% of TR energy belongs to an interval of  $0.1E_c$  to  $E_c$ . For large  $\omega_1$  the detector must have a larger thickness to absorb x-rays in this range. It would be important to control  $\omega_1$  during radiator production and use it as a free parameter in the detector optimization process.

Si-microstrip tracking detectors operating in a magnetic field can also be used for TR detection, even though the  $dE/dx$  losses in Si are much larger than the absorbed TR energy. The excellent spatial resolution of the Si detectors provides separation of the TR photons and  $dE/dx$  losses at relatively modest distances between radiator and detector. Simulations made on the basis of the beam-test data results has shown that in a magnetic field of 2 T and for the geometry of

the ATLAS Si-tracker proposed for sLHC, a rejection factor of  $> 30$  can be obtained for an electron efficiency above 90% over a particle momentum range 2-30 GeV/c (Brigida *et al.* in [113] and [114]). New detector techniques for TRDs are also under development and among them one should mention GasPixel detectors which allow to obtain a space point accuracy of  $< 30 \mu\text{m}$  and exploit all details of the particle tracks to highlight individual TR clusters in the gas (F. Harjes *et al.* in [114]). Thin films of heavy scintillators (V.V. Berdnikov *et al.* in [114]) might be very attractive in a combination with new radiators mentioned above.

**34.6.7. Resistive-plate chambers :** Written July 2015 by G. Aielli (U. Roma Tor Vergata).

The resistive-plate chamber (RPC) is a gaseous detector developed by R. Santonico and R. Cardarelli in the early 1980's [119]\*. Although its first purpose was to provide a competitive alternative to large scintillator counters, it was quickly recognized that it had relevant potential as a timing tracker due to the high space-time localization of the discharge. The RPC, as sketched in Fig. 34.17, is a large planar capacitor with two parallel high bulk resistivity electrode plates ( $10^9$ – $10^{13} \Omega\cdot\text{cm}$ ) separated by a set of insulating spacers. The spacers define a gap in the range from a few millimeters down to 0.1 mm with a precision of a few  $\mu\text{m}$ . The gap is filled with a suitable atmospheric-pressure gas mixture which serves as a target for ionizing radiation. Primary ionization for sub-millimeter gas gaps can be insufficient, thus multiple gaps can be combined to ensure an acceptable detection efficiency [121].



**Figure 34.17:** Schematic cross section of a generic single gap RPC.

The electrodes are most commonly made of high pressure phenolic-melaminic laminate (HPL), improperly referred to as "bakelite", or glass. A moderate electrode resistivity ( $\sim 10^5 \Omega/\square$ ) establishes a uniform electric field of several kV/mm across the gap, which initiates an electron avalanche following primary ionization. The above resistivity is low enough to ensure uniformity of the electric field, yet still transparent to fast signal transients from avalanches. This field configuration allows an excellent space-time localization of the signal. Due to the high electrode resistivity in RPCs, the electrode time constant is much longer than discharge processes. Therefore only the locally-stored electrostatic energy contributes to the discharge, which prevents the formation of sparks and leaves the rest of the detector field unaffected. The gas-facing surface of HPL electrodes are commonly coated with a resistive varnish (e.g.  $\sim \mu\text{m}$  layer of polymerized linseed oil) to achieve the necessary resistivity as well as to protect the electrode from discharge damage. As with other gaseous detectors, the gas mixture is optimized for each specific application. In general it needs to contain a component to quench UV photons, thus avoiding discharge propagation. An electronegative component controls the avalanche growth in case of very high electric fields [122,123]. To first order, each primary ionization in an RPC is exponentially amplified according to its distance from the anode. Therefore RPC signals span a large dynamic range, unlike gaseous detectors where ionization and amplification occur in separate regions (e.g. wire chambers or MPGDs). For increasingly stronger fields, the avalanche exponential growth progressively saturates to linear [124], and finally reaches a strongly-saturated "streamer" transition which exhausts all the locally-available energy [125]. The signal induced by the fast movement of the avalanche electrons is

\* The RPC was based on earlier work on a spark counter with one metallic and one high-resistivity plate [120].

isotropically distributed with respect to the field direction and present with equal but opposite amplitude on the two electrodes. A set of metallic readout electrodes (e.g. pads or strips) placed behind the resistive electrodes detect the charge pulse. This feature allows for 2D localization of the signal with uniform spatial resolution. Sensitivity to high-frequency electron avalanche signals over large RPC areas requires a correspondingly adequate Faraday cage and readout structure design. In particular, the front end electronics must be time-sensitive with a fast response and low noise, although these requirements are usually in competition [126].

**34.6.7.1. RPC types and applications:** RPCs are generally classified in two categories depending on the gas gap structure: single gap RPCs (described above) and multiple gap RPCs (typically referred as mRPCs or timing RPCs). While they are both based on the same principle they have different construction techniques, performance and limitations, making them suitable for different applications. Due to its simplicity and robustness, the single gap RPC is ideal for covering very large surfaces. Typical detector systems can have sensitive surface areas up to  $\sim 10^4 \text{ m}^2$ , with single module areas of a few  $\text{m}^2$ , and a space-time resolution down to  $\sim 0.4 \text{ ns} \times 100 \text{ } \mu\text{m}$  [127,128]. Typical applications are in muon systems (e.g. the muon trigger systems of the LHC experiments) or ground and underground based cosmic rays and neutrino arrays [129]. Moreover, single gap RPCs have recently found an application in tracking calorimetry [130]. The mRPC allows for smaller gas gap thicknesses while still maintaining a sufficient gaseous target. The most common version [131] consists of a stack of floating glass electrodes separated by monofilament (i.e. fishing) line, sandwiched between two external electrodes which provide the high-voltage bias. The floating glass electrodes assume a potential determined by the avalanche processes occurring between them. mRPCs have been largely used in TOF systems and in applications such as timing PET.

**34.6.7.2. Time and space resolution:** The RPC field configuration generates an avalanche which is strongly correlated in space and time to the original ionizing event. Space-time uncertainties generally arise from the statistical fluctuations of the ionization and multiplication processes, and from the characteristics of the readout and front-end electronics. The intrinsic signal latency is commonly in the ns range, making the RPC suitable for applications where a low latency is essential. A higher time resolution and shorter signal duration is correlated with a thinner gas gap, although a higher electric field is required for sufficient avalanche development [131,132]. Typical timing performances are from around 1 ns with a 2 mm gas gap, down to 20 ps for a stack of several 0.1 mm gaps [133]. The mechanical delicacy of sub-mm gap structures currently limit this technique to small detector areas. Digital strip readouts are commonly used, with spatial resolution determined by the strip pitch and the cluster size ( $\sim 0.5 \text{ cm}$ ). Recent developments toward higher spatial resolutions are mostly based on charge centroid techniques, benefiting from the availability of low-cost high-performance readout electronics. The present state of the art detectors have a combined space-time resolution of  $\sim 50 \text{ ps} \times 40 \text{ } \mu\text{m}$  [134].

**34.6.7.3. Rate capability and ageing:** RPC rate capability is limited by the voltage drop on resistive electrodes,  $\Delta V = V_a - V_{\text{gas}} = I \cdot R$  [135]. Here  $V_a$  is the applied voltage,  $V_{\text{gas}}$  is the effective voltage on the gas,  $R = \rho \cdot d/S$  is the total electrode resistance and  $I$  is the working current. Expressing  $I$  as the particle flux  $\Phi$  times an average charge per avalanche  $\langle Q \rangle$  gives  $\Delta V/\Phi = \rho \cdot d \cdot \langle Q \rangle$ . A large  $I$  not only limits the rate capability but also affects the long term performance of the detector. Discharges deplete the conductive properties of HPL electrodes [136]. In the presence of fluorocarbons and water, discharges generate hydrofluoric acid (HF) which damages internal detector surfaces, particularly glass electrodes [137]. HF damage can be mitigated by preventing water vapor contamination and by sufficient flushing of the gas gap. Operating in the streamer regime puts low requirements on the front end electronics sensitivity, but generally limits the counting rate capability to  $\sim 100 \text{ Hz/cm}^2$  and requires stability over a large gain range. Higher rate operation can be achieved by reducing gas gain in favor of electronic amplification. Increasing electronegative gases, such as  $\text{C}_2\text{H}_2\text{F}_4$  and  $\text{SF}_6$  [123], shifts

the streamer transition to higher gains. With these techniques, stable performance at high rates (e.g.  $10 \text{ kHz/cm}^2$ ) has been achieved for large area single gap RPCs [126]. Additional techniques rely on the natural redundancy and small gain of multiple gap structures [138] and electrodes made with lower resistivity materials [139].

### 34.7. Semiconductor detectors

Updated November 2013 by H. Spieler.

Semiconductor detectors provide a unique combination of energy and position resolution. In collider detectors they are most widely used as position sensing devices and photodetectors (Sec. 34.2). Integrated circuit technology allows the formation of high-density micron-scale electrodes on large (15–20 cm diameter) wafers, providing excellent position resolution. Furthermore, the density of silicon and its small ionization energy yield adequate signals with active layers only 100–300  $\mu\text{m}$  thick, so the signals are also fast (typically tens of ns). The high energy resolution is a key parameter in x-ray, gamma, and charged particle spectroscopy, e.g., in neutrinoless double beta decay searches. Silicon and germanium are the most commonly used materials, but gallium-arsenide, CdTe, CdZnTe, and other materials are also useful. CdZnTe provides a higher stopping power and the ratio of Cd to Zn concentrations changes the bandgap. Ge detectors are commonly operated at liquid nitrogen temperature to reduce the bias current, which depends exponentially on temperature. Semiconductor detectors depend crucially on low-noise electronics (see Sec. 34.8), so the detection sensitivity is determined by signal charge and capacitance. For a comprehensive discussion of semiconductor detectors and electronics see Ref. 140 or the tutorial website <http://www-physics.lbl.gov/spieler>.

#### 34.7.1. Materials Requirements :

Semiconductor detectors are essentially solid state ionization chambers. Absorbed energy forms electron-hole pairs, i.e., negative and positive charge carriers, which under an applied electric field move towards their respective collection electrodes, where they induce a signal current. The energy required to form an electron-hole pair is proportional to the bandgap. In tracking detectors the energy loss in the detector should be minimal, whereas for energy spectroscopy the stopping power should be maximized, so for gamma rays high- $Z$  materials are desirable.

Measurements on silicon photodiodes [141] show that for photon energies below 4 eV one electron-hole ( $e-h$ ) pair is formed per incident photon. The mean energy  $E_i$  required to produce an  $e-h$  pair peaks at 4.4 eV for a photon energy around 6 eV. Above  $\sim 1.5 \text{ keV}$  it assumes a constant value, 3.67 eV at room temperature. It is larger than the bandgap energy because momentum conservation requires excitation of lattice vibrations (phonons). For minimum-ionizing particles, the most probable charge deposition in a 300  $\mu\text{m}$  thick silicon detector is about 3.5 fC (22000 electrons). Other typical ionization energies are 2.96 eV in Ge, 4.2 eV in GaAs, and 4.43 eV in CdTe.

Since both electronic and lattice excitations are involved, the variance in the number of charge carriers  $N = E/E_i$  produced by an absorbed energy  $E$  is reduced by the Fano factor  $F$  (about 0.1 in Si and Ge). Thus,  $\sigma_N = \sqrt{FN}$  and the energy resolution  $\sigma_E/E = \sqrt{FE_i/E}$ . However, the measured signal fluctuations are usually dominated by electronic noise or energy loss fluctuations in the detector. The electronic noise contributions depend on the pulse shaping in the signal processing electronics, so the choice of the shaping time is critical (see Sec. 34.8).

A smaller bandgap would produce a larger signal and improve energy resolution, but the intrinsic resistance of the material is critical. Thermal excitation, given by the Fermi-Dirac distribution, promotes electrons into the conduction band, so the thermally excited carrier concentration increases exponentially with decreasing bandgaps. In pure Si the carrier concentration is  $\sim 10^{10} \text{ cm}^{-3}$  at 300 K, corresponding to a resistivity  $\rho \approx 400 \text{ k}\Omega\text{cm}$ . In reality, crystal imperfections and minute impurity concentrations limit Si carrier concentrations to  $\sim 10^{11} \text{ cm}^{-3}$  at 300 K, corresponding to a resistivity  $\rho \approx 40 \text{ k}\Omega\text{cm}$ . In practice, resistivities up to 20 k $\Omega\text{cm}$  are available, with mass production ranging from 5 to 10 k $\Omega\text{cm}$ . Signal currents at keV scale

energies are of order  $\mu\text{A}$ . However, for a resistivity of  $10^4 \Omega\text{cm}$  a  $300 \mu\text{m}$  thick sensor with  $1 \text{ cm}^2$  area would have a resistance of  $300 \Omega$ , so  $30 \text{ V}$  would lead to a current flow of  $100 \text{ mA}$  and a power dissipation of  $3 \text{ W}$ . On the other hand, high-quality single crystals of Si and Ge can be grown economically with suitably large volumes, so to mitigate the effect of resistivity one resorts to reverse-biased diode structures. Although this reduces the bias current relative to a resistive material, the thermally excited leakage current can still be excessive at room temperature, so Ge diodes are typically operated at liquid nitrogen temperature ( $77 \text{ K}$ ).

A major effort is to find high- $Z$  materials with a bandgap that is sufficiently high to allow room-temperature operation while still providing good energy resolution. Compound semiconductors, *e.g.*, CdZnTe, can allow this, but typically suffer from charge collection problems, characterized by the product  $\mu\tau$  of mobility and carrier lifetime. In Si and Ge  $\mu\tau > 1 \text{ cm}^2 \text{ V}^{-1}$  for both electrons and holes, whereas in compound semiconductors it is in the range  $10^{-3}$ – $10^{-8}$ . Since for holes  $\mu\tau$  is typically an order of magnitude smaller than for electrons, detector configurations where the electron contribution to the charge signal dominates—*e.g.*, strip or pixel structures—can provide better performance.

### 34.7.2. Detector Configurations :

A  $p$ - $n$  junction operated at reverse bias forms a sensitive region depleted of mobile charge and sets up an electric field that sweeps charge liberated by radiation to the electrodes. Detectors typically use an asymmetric structure, *e.g.*, a highly doped  $p$  electrode and a lightly doped  $n$  region, so that the depletion region extends predominantly into the lightly doped volume.

In a planar device the thickness of the depleted region is

$$W = \sqrt{2\epsilon(V + V_{bi})/Ne} = \sqrt{2\rho\mu\epsilon(V + V_{bi})}, \quad (34.19)$$

where  $V$  = external bias voltage

$V_{bi}$  = “built-in” voltage ( $\approx 0.5 \text{ V}$  for resistivities typically used in Si detectors)

$N$  = doping concentration

$e$  = electronic charge

$\epsilon$  = dielectric constant =  $11.9 \epsilon_0 \approx 1 \text{ pF/cm}$  in Si

$\rho$  = resistivity (typically  $1$ – $10 \text{ k}\Omega \text{ cm}$  in Si)

$\mu$  = charge carrier mobility

=  $1350 \text{ cm}^2 \text{ V}^{-1} \text{ s}^{-1}$  for electrons in Si

=  $450 \text{ cm}^2 \text{ V}^{-1} \text{ s}^{-1}$  for holes in Si

In Si

$$W = 0.5 [\mu\text{m}/\sqrt{\Omega\text{-cm} \cdot \text{V}}] \times \sqrt{\rho(V + V_{bi})} \text{ for } n\text{-type Si, and}$$

$$W = 0.3 [\mu\text{m}/\sqrt{\Omega\text{-cm} \cdot \text{V}}] \times \sqrt{\rho(V + V_{bi})} \text{ for } p\text{-type Si.}$$

The conductive  $p$  and  $n$  regions together with the depleted volume form a capacitor with the capacitance per unit area

$$C = \epsilon/W \approx 1 [\text{pF/cm}] / W \text{ in Si.} \quad (34.20)$$

In strip and pixel detectors the capacitance is dominated by the fringing capacitance to neighboring electrodes. For example, the strip-to-strip Si fringing capacitance is  $\sim 1$ – $1.5 \text{ pF cm}^{-1}$  of strip length at a strip pitch of  $25$ – $50 \mu\text{m}$ .

Large volume ( $\sim 10^2$ – $10^3 \text{ cm}^3$ ) Ge detectors are commonly configured as coaxial detectors, *e.g.*, a cylindrical  $n$ -type crystal with  $5$ – $10 \text{ cm}$  diameter and  $10 \text{ cm}$  length with an inner  $5$ – $10 \text{ mm}$  diameter  $n^+$  electrode and an outer  $p^+$  layer forming the diode junction. Ge can be grown with very low impurity levels,  $10^9$ – $10^{10} \text{ cm}^{-3}$  (HPGe), so these large volumes can be depleted with several kV.

### 34.7.3. Signal Formation :

The signal pulse shape depends on the instantaneous carrier velocity  $v(x) = \mu E(x)$  and the electrode geometry, which determines the distribution of induced charge (*e.g.*, see Ref. 140, pp. 71–83). Charge collection time decreases with increasing bias voltage, and can be reduced further by operating the detector with “overbias,” *i.e.*, a bias voltage exceeding the value required to fully deplete the device. Note that in partial depletion the electric field goes to zero, whereas going beyond full depletion adds a constantly distributed field. The collection time is limited by velocity saturation at high fields (in Si approaching  $10^7 \text{ cm/s}$  at  $E > 10^4 \text{ V/cm}$ ); at an average field of  $10^4 \text{ V/cm}$  the collection time is about  $15 \text{ ps}/\mu\text{m}$  for electrons and  $30 \text{ ps}/\mu\text{m}$  for holes. In typical fully-depleted detectors  $300 \mu\text{m}$  thick, electrons are collected within about  $10 \text{ ns}$ , and holes within about  $25 \text{ ns}$ .

Position resolution is limited by transverse diffusion during charge collection (typically  $5 \mu\text{m}$  for  $300 \mu\text{m}$  thickness) and by knock-on electrons. Resolutions of  $2$ – $4 \mu\text{m}$  (rms) have been obtained in beam tests. In magnetic fields, the Lorentz drift deflects the electron and hole trajectories and the detector must be tilted to reduce spatial spreading (see “Hall effect” in semiconductor textbooks).

Electrodes can be in the form of cm-scale pads, strips, or  $\mu\text{m}$ -scale pixels. Various readout structures have been developed for pixels, *e.g.*, CCDs, DEPFETs, monolithic pixel devices that integrate sensor and electronics (MAPS), and hybrid pixel devices that utilize separate sensors and readout ICs connected by two-dimensional arrays of solder bumps. For an overview and further discussion see Ref. 140.

In gamma ray spectroscopy ( $E_\gamma > 10^2 \text{ keV}$ ) Compton scattering dominates, so for a significant fraction of events the incident gamma energy is not completely absorbed, *i.e.*, the Compton scattered photon escapes from the detector and the energy deposited by the Compton electron is only a fraction of the total. Distinguishing multi-interaction events, *e.g.*, multiple Compton scatters with a final photoelectric absorption, from single Compton scatters allows background suppression. Since the individual interactions take place in different parts of the detector volume, these events can be distinguished by segmenting the outer electrode of a coaxial detector and analyzing the current pulse shapes. The different collection times can be made more distinguishable by using “point” electrodes, where most of the signal is induced when charges are close to the electrode, similarly to strip or pixel detectors. Charge clusters arriving from different positions in the detector will arrive at different times and produce current pulses whose major components are separated in time. Point electrodes also reduce the electrode capacitance, which reduces electronic noise, but careful design is necessary to avoid low-field regions in the detector volume.

**34.7.4. Radiation Damage :** Radiation damage occurs through two basic mechanisms:

1. Bulk damage due to displacement of atoms from their lattice sites. This leads to increased leakage current, carrier trapping, and build-up of space charge that changes the required operating voltage. Displacement damage depends on the nonionizing energy loss and the energy imparted to the recoil atoms, which can initiate a chain of subsequent displacements, *i.e.*, damage clusters. Hence, it is critical to consider both particle type and energy.
2. Surface damage due to charge build-up in surface layers, which leads to increased surface leakage currents. In strip detectors the inter-strip isolation is affected. The effects of charge build-up are strongly dependent on the device structure and on fabrication details. Since the damage is proportional to the absorbed energy (when ionization dominates), the dose can be specified in rad (or Gray) independent of particle type.

The increase in reverse bias current due to bulk damage is  $\Delta I_r = \alpha \Phi$  per unit volume, where  $\Phi$  is the particle fluence and  $\alpha$  the damage coefficient ( $\alpha \approx 3 \times 10^{-17} \text{ A/cm}$  for minimum ionizing protons and pions after long-term annealing;  $\alpha \approx 2 \times 10^{-17} \text{ A/cm}$  for  $1 \text{ MeV}$

neutrons). The reverse bias current depends strongly on temperature

$$\frac{I_R(T_2)}{I_R(T_1)} = \left(\frac{T_2}{T_1}\right)^2 \exp\left[-\frac{E}{2k}\left(\frac{T_1 - T_2}{T_1 T_2}\right)\right], \quad (34.21)$$

where  $E = 1.2$  eV, so rather modest cooling can reduce the current substantially ( $\sim 6$ -fold current reduction in cooling from room temperature to  $0^\circ\text{C}$ ).

Displacement damage forms acceptor-like states. These trap electrons, building up a negative space charge, which in turn requires an increase in the applied voltage to sweep signal charge through the detector thickness. This has the same effect as a change in resistivity, *i.e.*, the required voltage drops initially with fluence, until the positive and negative space charge balance and very little voltage is required to collect all signal charge. At larger fluences the negative space charge dominates, and the required operating voltage increases ( $V \propto N$ ). The safe limit on operating voltage ultimately limits the detector lifetime. Strip detectors specifically designed for high voltages have been extensively operated at bias voltages  $>500$  V. Since the effect of radiation damage depends on the electronic activity of defects, various techniques have been applied to neutralize the damage sites. For example, additional doping with oxygen can increase the allowable charged hadron fluence roughly three-fold [142]. Detectors with columnar electrodes normal to the surface can also extend operational lifetime [143]. The increase in leakage current with fluence, on the other hand, appears to be unaffected by resistivity and whether the material is *n* or *p*-type. At fluences beyond  $10^{15} \text{ cm}^{-2}$  decreased carrier lifetime becomes critical [144,145].

Strip and pixel detectors have remained functional at fluences beyond  $10^{15} \text{ cm}^{-2}$  for minimum ionizing protons. At this damage level, charge loss due to recombination and trapping becomes significant and the high signal-to-noise ratio obtainable with low-capacitance pixel structures extends detector lifetime. The higher mobility of electrons makes them less sensitive to carrier lifetime than holes, so detector configurations that emphasize the electron contribution to the charge signal are advantageous, *e.g.*,  $n^+$  strips or pixels on a *p*- or *n*-substrate. The occupancy of the defect charge states is strongly temperature dependent; competing processes can increase or decrease the required operating voltage. It is critical to choose the operating temperature judiciously ( $-10$  to  $0^\circ\text{C}$  in typical collider detectors) and limit warm-up periods during maintenance. For a more detailed summary see Ref. 146 and the web-sites of the ROSE and RD50 collaborations at <http://RD48.web.cern.ch/rd48> and <http://RD50.web.cern.ch/rd50>. Materials engineering, *e.g.*, introducing oxygen interstitials, can improve certain aspects and is under investigation. At high fluences diamond is an alternative, but operates as an insulator rather than a reverse-biased diode.

Currently, the lifetime of detector systems is still limited by the detectors; in the electronics use of standard “deep submicron” CMOS fabrication processes with appropriately designed circuitry has increased the radiation resistance to fluences  $> 10^{15} \text{ cm}^{-2}$  of minimum ionizing protons or pions. For a comprehensive discussion of radiation effects see Ref. 147.

### 34.8. Low-noise electronics

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Many detectors rely critically on low-noise electronics, either to improve energy resolution or to allow a low detection threshold. A typical detector front-end is shown in Fig. 34.18.

The detector is represented by a capacitance  $C_d$ , a relevant model for most detectors. Bias voltage is applied through resistor  $R_b$  and the signal is coupled to the preamplifier through a blocking capacitor  $C_c$ . The series resistance  $R_s$  represents the sum of all resistances present in the input signal path, *e.g.* the electrode resistance, any input protection networks, and parasitic resistances in the input transistor. The preamplifier provides gain and feeds a pulse shaper, which tailors the overall frequency response to optimize signal-to-noise ratio while limiting the duration of the signal pulse to accommodate the signal pulse rate. Even if not explicitly stated, all amplifiers provide some form of pulse shaping due to their limited frequency response.

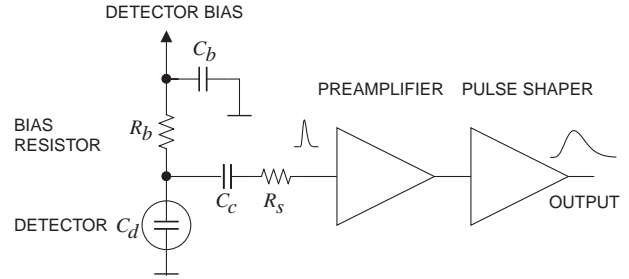


Figure 34.18: Typical detector front-end circuit.

The equivalent circuit for the noise analysis (Fig. 34.19) includes both current and voltage noise sources. The leakage current of a semiconductor detector, for example, fluctuates due to continuous electron emission statistics. The statistical fluctuations in the charge measurement will scale with the square root of the total number of recorded charges, so this noise contribution increases with the width of the shaped output pulse. This “shot noise”  $i_{nd}$  is represented by a current noise generator in parallel with the detector. Resistors exhibit noise due to thermal velocity fluctuations of the charge carriers. This yields a constant noise power density vs. frequency, so increasing the bandwidth of the shaped output pulse, *i.e.* reducing the shaping time, will increase the noise. This noise source can be modeled either as a voltage or current generator. Generally, resistors shunting the input act as noise current sources and resistors in series with the input act as noise voltage sources (which is why some in the detector community refer to current and voltage noise as “parallel” and “series” noise). Since the bias resistor effectively shunts the input, as the capacitor  $C_b$  passes current fluctuations to ground, it acts as a current generator  $i_{nb}$  and its noise current has the same effect as the shot noise current from the detector. Any other shunt resistances can be incorporated in the same way. Conversely, the series resistor  $R_s$  acts as a voltage generator. The electronic noise of the amplifier is described fully by a combination of voltage and current sources at its input, shown as  $e_{na}$  and  $i_{na}$ .

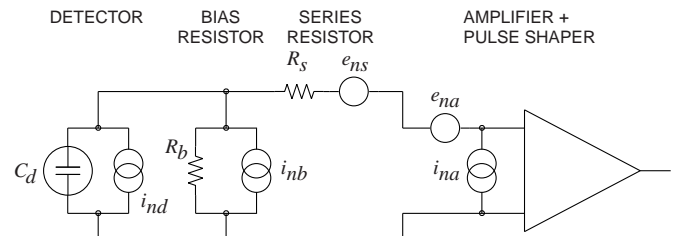


Figure 34.19: Equivalent circuit for noise analysis.

Shot noise and thermal noise have a “white” frequency distribution, *i.e.* the spectral power densities  $dP_n/df \propto di_n^2/df \propto de_n^2/df$  are constant with the magnitudes

$$\begin{aligned} i_{nd}^2 &= 2eI_d, \\ i_{nb}^2 &= \frac{4kT}{R_b}, \\ e_{ns}^2 &= 4kTR_s, \end{aligned} \quad (34.22)$$

where  $e$  is the electronic charge,  $I_d$  the detector bias current,  $k$  the Boltzmann constant and  $T$  the temperature. Typical amplifier noise parameters  $e_{na}$  and  $i_{na}$  are of order  $\text{nV}/\sqrt{\text{Hz}}$  and  $\text{pA}/\sqrt{\text{Hz}}$ . Trapping and detrapping processes in resistors, dielectrics and semiconductors can introduce additional fluctuations whose noise power frequently exhibits a  $1/f$  spectrum. The spectral density of the  $1/f$  noise voltage is

$$e_{nf}^2 = \frac{A_f}{f}, \quad (34.23)$$

where the noise coefficient  $A_f$  is device specific and of order  $10^{-10}$ – $10^{-12} \text{ V}^2$ .

A fraction of the noise current flows through the detector capacitance, resulting in a frequency-dependent noise voltage  $i_n/(\omega C_d)$ , which is added to the noise voltage in the input circuit. Thus, the current noise contribution increases with lowering frequency, so its contribution increases with shaping pulse width. Since the individual noise contributions are random and uncorrelated, they add in quadrature. The total noise at the output of the pulse shaper is obtained by integrating over the full bandwidth of the system. Superimposed on repetitive detector signal pulses of constant magnitude, purely random noise produces a Gaussian signal distribution.

Since radiation detectors typically convert the deposited energy into charge, the system's noise level is conveniently expressed as an equivalent noise charge  $Q_n$ , which is equal to the detector signal that yields a signal-to-noise ratio of one. The equivalent noise charge is commonly expressed in Coulombs, the corresponding number of electrons, or the equivalent deposited energy (eV). For a capacitive sensor

$$Q_n^2 = i_n^2 F_i T_S + e_n^2 F_v \frac{C^2}{T_S} + F_v f A_f C^2, \quad (34.24)$$

where  $C$  is the sum of all capacitances shunting the input,  $F_i$ ,  $F_v$ , and  $F_v f$  depend on the shape of the pulse determined by the shaper and  $T_S$  is a characteristic time, for example, the peaking time of a semi-gaussian pulse or the sampling interval in a correlated double sampler. The form factors  $F_i$ ,  $F_v$  are easily calculated

$$F_i = \frac{1}{2T_S} \int_{-\infty}^{\infty} [W(t)]^2 dt, \quad F_v = \frac{T_S}{2} \int_{-\infty}^{\infty} \left[ \frac{dW(t)}{dt} \right]^2 dt, \quad (34.25)$$

where for time-invariant pulse-shaping  $W(t)$  is simply the system's impulse response (the output signal seen on an oscilloscope) for a short input pulse with the peak output signal normalized to unity. For more details see Refs. 148 and 149.

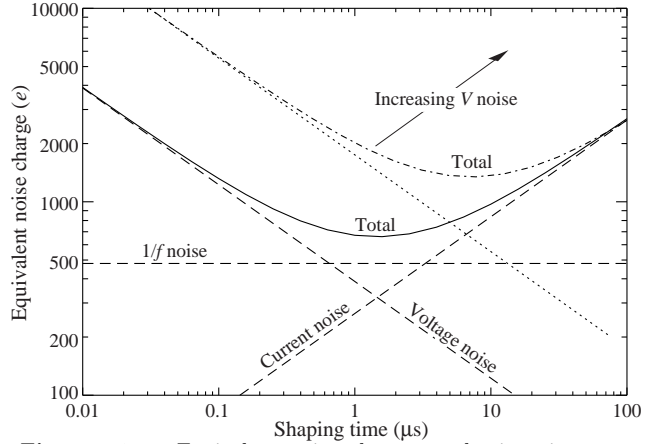
A pulse shaper formed by a single differentiator and integrator with equal time constants has  $F_i = F_v = 0.9$  and  $F_v f = 4$ , independent of the shaping time constant. The overall noise bandwidth, however, depends on the time constant, *i.e.* the characteristic time  $T_S$ . The contribution from noise currents increases with shaping time, *i.e.*, pulse duration, whereas the voltage noise decreases with increasing shaping time, *i.e.* reduced bandwidth. Noise with a  $1/f$  spectrum depends only on the ratio of upper to lower cutoff frequencies (integrator to differentiator time constants), so for a given shaper topology the  $1/f$  contribution to  $Q_n$  is independent of  $T_S$ . Furthermore, the contribution of noise voltage sources to  $Q_n$  increases with detector capacitance. Pulse shapers can be designed to reduce the effect of current noise, *e.g.*, mitigate radiation damage. Increasing pulse symmetry tends to decrease  $F_i$  and increase  $F_v$  (*e.g.*, to 0.45 and 1.0 for a shaper with one  $CR$  differentiator and four cascaded integrators). For the circuit shown in Fig. 34.19,

$$Q_n^2 = (2eI_d + 4kT/R_b + i_{na}^2) F_i T_S + (4kTR_s + e_{na}^2) F_v C_d^2 / T_S + F_v f A_f C_d^2. \quad (34.26)$$

As the characteristic time  $T_S$  is changed, the total noise goes through a minimum, where the current and voltage contributions are equal. Fig. 34.20 shows a typical example. At short shaping times the voltage noise dominates, whereas at long shaping times the current noise takes over. The noise minimum is flattened by the presence of  $1/f$  noise. Increasing the detector capacitance will increase the voltage noise and shift the noise minimum to longer shaping times.

For quick estimates, one can use the following equation, which assumes an FET amplifier (negligible  $i_{na}$ ) and a simple  $CR$ - $RC$  shaper with time constants  $\tau$  (equal to the peaking time):

$$(Q_n/e)^2 = 12 \left[ \frac{1}{\text{nA} \cdot \text{ns}} \right] I_d \tau + 6 \times 10^5 \left[ \frac{\text{k}\Omega}{\text{ns}} \right] \frac{\tau}{R_b} + 3.6 \times 10^4 \left[ \frac{\text{ns}}{(\text{pF})^2 (\text{nV})^2 / \text{Hz}} \right] e_n^2 \frac{C^2}{\tau}. \quad (34.27)$$



**Figure 34.20:** Equivalent noise charge *vs* shaping time. Changing the voltage or current noise contribution shifts the noise minimum. Increased voltage noise is shown as an example.

Noise is improved by reducing the detector capacitance and leakage current, judiciously selecting all resistances in the input circuit, and choosing the optimum shaping time constant. Another noise contribution to consider is that noise cross-couples from the neighboring front-ends in strip and pixel detectors through the inter-electrode capacitance.

The noise parameters of the amplifier depend primarily on the input device. In field effect transistors, the noise current contribution is very small, so reducing the detector leakage current and increasing the bias resistance will allow long shaping times with correspondingly lower noise. In bipolar transistors, the base current sets a lower bound on the noise current, so these devices are best at short shaping times. In special cases where the noise of a transistor scales with geometry, *i.e.*, decreasing noise voltage with increasing input capacitance, the lowest noise is obtained when the input capacitance of the transistor is equal to the detector capacitance, albeit at the expense of power dissipation. Capacitive matching is useful with field-effect transistors, but not bipolar transistors. In bipolar transistors, the minimum obtainable noise is independent of shaping time, but only at the optimum collector current  $I_C$ , which does depend on shaping time.

$$Q_{n,\min}^2 = 4kT \frac{C}{\sqrt{\beta_{DC}}} \sqrt{F_i F_v} \quad \text{at} \quad I_C = \frac{kT}{e} C \sqrt{\beta_{DC}} \sqrt{\frac{F_v}{F_i} \frac{1}{T_S}}, \quad (34.28)$$

where  $\beta_{DC}$  is the DC current gain. For a  $CR$ - $RC$  shaper and  $\beta_{DC} = 100$ ,

$$Q_{n,\min}/e \approx 250 \sqrt{C/\text{pF}}. \quad (34.29)$$

Practical noise levels range from  $\sim 1e$  for CCD's at long shaping times to  $\sim 10^4 e$  in high-capacitance liquid argon calorimeters. Silicon strip detectors typically operate at  $\sim 10^3$  electrons, whereas pixel detectors with fast readout provide noise of several hundred electrons.

In timing measurements, the slope-to-noise ratio must be optimized, rather than the signal-to-noise ratio alone, so the rise time  $t_r$  of the pulse is important. The "jitter"  $\sigma_t$  of the timing distribution is

$$\sigma_t = \frac{\sigma_n}{(dS/dt)_{S_T}} \approx \frac{t_r}{S/N}, \quad (34.30)$$

where  $\sigma_n$  is the rms noise and the derivative of the signal  $dS/dt$  is evaluated at the trigger level  $S_T$ . To increase  $dS/dt$  without incurring excessive noise, the amplifier bandwidth should match the rise-time of the detector signal. The 10 to 90% rise time of an amplifier with bandwidth  $f_U$  is  $0.35/f_U$ . For example, an oscilloscope with 350 MHz bandwidth has a 1 ns rise time. When amplifiers are cascaded, which is invariably necessary, the individual rise times add in quadrature.

$$t_r \approx \sqrt{t_{r1}^2 + t_{r2}^2 + \dots + t_{rn}^2}. \quad (34.31)$$



Increasing signal-to-noise ratio also improves time resolution, so minimizing the total capacitance at the input is also important. At high signal-to-noise ratios, the time jitter can be much smaller than the rise time. The timing distribution may shift with signal level (“walk”), but this can be corrected by various means, either in hardware or software [8].

The basic principles discussed above apply to both analog and digital signal processing. In digital signal processing the pulse shaper shown in Fig. 34.18 is replaced by an analog to digital converter (ADC) followed by a digital processor that determines the pulse shape. Digital signal processing allows great flexibility in implementing filtering functions. The software can be changed readily to adapt to a wide variety of operating conditions and it is possible to implement filters that are impractical or even impossible using analog circuitry. However, this comes at the expense of increased circuit complexity and increased demands on the ADC compared to analog shaping.

If the sampling rate of the ADC is too low, high frequency components will be transferred to lower frequencies (“aliasing”). The sampling rate of the ADC must be high enough to capture the maximum frequency component of the input signal. Apart from missing information on the fast components of the pulse, undersampling introduces spurious artifacts. If the frequency range of the input signal is much greater, the noise at the higher frequencies will be transferred to lower frequencies and increase the noise level in the frequency range of pulses formed in the subsequent digital shaper. The Nyquist criterion states that the sampling frequency must be at least twice the maximum relevant input frequency. This requires that the bandwidth of the circuitry preceding the ADC must be limited. The most reliable technique is to insert a low-pass filter.

The digitization process also introduces inherent noise, since the voltage range  $\Delta V$  corresponding to a minimum bit introduces quasi-random fluctuations relative to the exact amplitude

$$\sigma_n = \frac{\Delta V}{\sqrt{12}}. \quad (34.32)$$

When the Nyquist condition is fulfilled the noise bandwidth  $\Delta f_n$  is spread nearly uniformly and extends to 1/2 the sampling frequency  $f_S$ , so the spectral noise density

$$e_n = \frac{\sigma_n}{\sqrt{\Delta f_n}} = \frac{\Delta V}{\sqrt{12}} \cdot \frac{1}{\sqrt{f_S/2}} = \frac{\Delta V}{\sqrt{6}f_S}. \quad (34.33)$$

Sampling at a higher frequency spreads the total noise over a larger frequency range, so oversampling can be used to increase the effective resolution. In practice, this quantization noise is increased by differential nonlinearity. Furthermore, the equivalent input noise of ADCs is often rather high, so the overall gain of the stages preceding the ADC must be sufficiently large for the preamplifier input noise to override.

When implemented properly, digital signal processing provides significant advantages in systems where the shape of detector signal pulses changes greatly, for example in large semiconductor detectors for gamma rays or in gaseous detectors (*e.g.* TPCs) where the duration of the current pulse varies with drift time, which can range over orders of magnitude. Where is analog signal processing best (most efficient)? In systems that require fast time response the high power requirements of high-speed ADCs are prohibitive. Systems that are not sensitive to pulse shape can use fixed shaper constants and rather simple filters, which can be either continuous or sampled. In high density systems that require small circuit area and low power (*e.g.* strip and pixel detectors), analog filtering often yields the required response and tends to be most efficient.

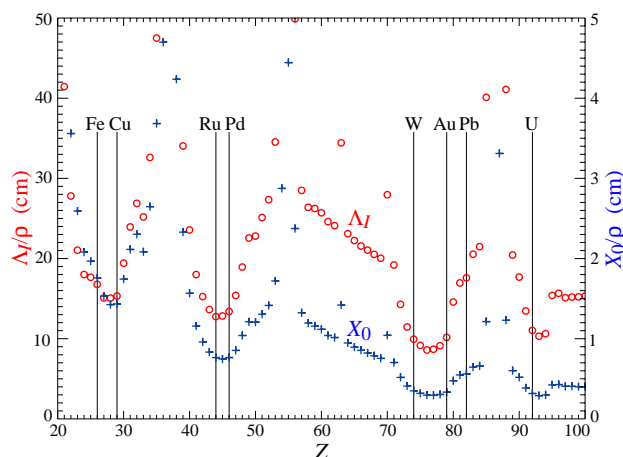
It is important to consider that additional noise is often introduced by external electronics, *e.g.* power supplies and digital systems. External noise can couple to the input. Often the “common grounding” allows additional noise current to couple to the current loop connecting the detector to the preamp. Recognizing additional noise sources and minimizing cross-coupling to the detector current loop is often important. Understanding basic physics and its practical effects is important in forming a broad view of the detector system

and recognizing potential problems (*e.g.* modified data), rather than merely following standard recipes.

For a more detailed introduction to detector signal processing and electronics see Ref. 140 or the tutorial website <http://www-physics.lbl.gov/spieler>.

### 34.9. Calorimeters

A calorimeter is designed to measure a particle’s (or jet’s) energy and direction for an (ideally) contained electromagnetic (EM) or hadronic shower. The characteristic interaction distance for an electromagnetic interaction is the radiation length  $X_0$ , which ranges from 13.8 g cm<sup>-2</sup> in iron to 6.0 g cm<sup>-2</sup> in uranium.\* Similarly, the characteristic nuclear interaction length  $\lambda_I$  varies from 132.1 g cm<sup>-2</sup> (Fe) to 209 g cm<sup>-2</sup> (U).† In either case, a calorimeter must be many interaction lengths deep, where “many” is determined by physical size, cost, and other factors. EM calorimeters tend to be 15–30  $X_0$  deep, while hadronic calorimeters are usually compromised at 5–8  $\lambda_I$ . In real experiments there is likely to be an EM calorimeter in front of the hadronic section, which in turn has less sampling density in the back, so the hadronic cascade occurs in a succession of different structures.



**Figure 34.21:** Nuclear interaction length  $\lambda_I/\rho$  (circles) and radiation length  $X_0/\rho$  (+s) in cm for the chemical elements with  $Z > 20$  and  $\lambda_I < 50$  cm.

In all cases there is a premium on small  $\lambda_I/\rho$  and  $X_0/\rho$  (both with units of length). These quantities are shown for  $Z > 20$  for the chemical elements in Fig. 34.21. For the hadronic case, metallic absorbers in the W–Au region are best, followed by U. The Ru–Pd region elements are rare and expensive. Lead is a bad choice. Given cost considerations, Fe and Cu might be appropriate choices. For EM calorimeters high  $Z$  is preferred, and lead is not a bad choice.

These considerations are for *sampling calorimeters* consisting of metallic absorber sandwiched or (threaded) with an active material which generates signal. The active medium may be a scintillator, an ionizing noble liquid, a gas chamber, a semiconductor, or a Cherenkov radiator. The average interaction length is thus greater than that of the absorber alone, sometimes substantially so.

There are also *homogeneous calorimeters*, in which the entire volume is sensitive, *i.e.*, contributes signal. Homogeneous calorimeters (so far usually electromagnetic) may be built with inorganic heavy (high density, high  $\langle Z \rangle$ ) scintillating crystals, or non-scintillating Cherenkov radiators such as lead glass and lead fluoride. Scintillation light and/or ionization in noble liquids can be detected. Nuclear interaction lengths in inorganic crystals range from 17.8 cm (LuAlO<sub>3</sub>) to 42.2 cm (NaI). Popular choices have been BGO with  $\lambda_I = 22.3$  cm and  $X_0 = 1.12$  cm, and PbWO<sub>4</sub> (20.3 cm and 0.89 cm). Properties of these and other commonly used inorganic crystal scintillators can be found in Table 34.4.

\*  $X_0 = 120 \text{ g cm}^{-2} Z^{-2/3}$  to better than 5% for  $Z > 23$ .

†  $\lambda_I = 37.8 \text{ g cm}^{-2} A^{0.312}$  to within 0.8% for  $Z > 15$ .

See [pdg.lbl.gov/AtomicNuclearProperties](http://pdg.lbl.gov/AtomicNuclearProperties) for actual values.

### 34.9.1. Electromagnetic calorimeters :

Revised September 2015 by R.-Y. Zhu (California Institute of Technology).

The development of electromagnetic showers is discussed in the section on “Passage of Particles Through Matter” (Sec. 33 of this *Review*). Formulae are given which approximately describe average showers, but since the physics of electromagnetic showers is well understood, detailed and reliable Monte Carlo simulation is possible. EGS4 [150] and GEANT [151] have emerged as the standards.

There are homogeneous and sampling electromagnetic calorimeters. In a homogeneous calorimeter the entire volume is sensitive, *i.e.*, contributes signal. Homogeneous electromagnetic calorimeters may be built with inorganic heavy (high- $Z$ ) scintillating crystals such as BaF<sub>2</sub>, BGO, CsI, LYSO, NaI and PWO, non-scintillating Cherenkov radiators such as lead glass and lead fluoride (PbF<sub>2</sub>), or ionizing noble liquids. Properties of commonly used inorganic crystal scintillators can be found in Table 34.4. A sampling calorimeter consists of an active medium which generates signal and a passive medium which functions as an absorber. The active medium may be a scintillator, an ionizing noble liquid, a semiconductor, or a gas chamber. The passive medium is usually a material of high density, such as lead, tungsten, iron, copper, or depleted uranium.

The energy resolution  $\sigma_E/E$  of a calorimeter can be parameterized as  $a/\sqrt{E} \oplus b \oplus c/E$ , where  $\oplus$  represents addition in quadrature and  $E$  is in GeV. The stochastic term  $a$  represents statistics-related fluctuations such as intrinsic shower fluctuations, photoelectron statistics, dead material at the front of the calorimeter, and sampling fluctuations. For a fixed number of radiation lengths, the stochastic term  $a$  for a sampling calorimeter is expected to be proportional to  $\sqrt{t/f}$ , where  $t$  is plate thickness and  $f$  is sampling fraction [152,153]. While  $a$  is at a few percent level for a homogeneous calorimeter, it is typically 10% for sampling calorimeters.

The main contributions to the systematic, or constant, term  $b$  are detector non-uniformity and calibration uncertainty. In the case of the hadronic cascades discussed below, non-compensation also contributes to the constant term. One additional contribution to the constant term for calorimeters built for modern high-energy physics experiments, operated in a high-beam intensity environment, is radiation damage of the active medium. This can be mitigated by developing radiation-hard active media [51], by reducing the signal path length [52] and by frequent *in situ* calibration and monitoring [50,153]. With effort, the constant term  $b$  can be reduced to below one percent. The term  $c$  is due to electronic noise summed over readout channels within a few Molière radii. The best energy resolution for electromagnetic shower measurement is obtained in total absorption homogeneous calorimeters, *e.g.* calorimeters built with heavy crystal scintillators. These are used when ultimate performance is pursued.

The position resolution depends on the effective Molière radius and the transverse granularity of the calorimeter. Like the energy resolution, it can be factored as  $a/\sqrt{E} \oplus b$ , where  $a$  is a few to 20 mm and  $b$  can be as small as a fraction of mm for a dense calorimeter with fine granularity. Electromagnetic calorimeters may also provide direction measurement for electrons and photons. This is important for photon-related physics when there are uncertainties in event origin, since photons do not leave information in the particle tracking system. Typical photon angular resolution is about 45 mrad/ $\sqrt{E}$ , which can be provided by implementing longitudinal segmentation [154] for a sampling calorimeter or by adding a preshower detector [155] for a homogeneous calorimeter without longitudinal segmentation.

Novel technologies have been developed for electromagnetic calorimetry. New heavy crystal scintillators, such as PWO and LYSO:Ce (see Sec. 34.4), have attracted much attention. In some cases, such as PWO, it has received broad applications in high-energy and nuclear physics experiments. The “spaghetti” structure has been developed for sampling calorimetry with scintillating fibers as the sensitive medium. The “shashlik” structure has been developed for sampling calorimetry with wavelength shifting fibers functioning as both the converter and transporter for light generated in the sensitive medium. The “accordion” structure has been developed for sampling

calorimetry with ionizing noble liquid as the sensitive medium.

Table 34.8 provides a brief description of typical electromagnetic calorimeters built recently for high-energy physics experiments. Also listed in this table are calorimeter depths in radiation lengths ( $X_0$ ) and the achieved energy resolution. Whenever possible, the performance of calorimeters *in situ* is quoted, which is usually in good agreement with prototype test beam results as well as EGS or GEANT simulations, provided that all systematic effects are properly included. Detailed references on detector design and performance can be found in Appendix C of reference [153] and Proceedings of the International Conference series on Calorimetry in High Energy Physics.

**Table 34.8:** Resolution of typical electromagnetic calorimeters.  
 $E$  is in GeV.

Technology (Experiment)	Depth	Energy resolution	Date
NaI(Tl) (Crystal Ball)	20 $X_0$	2.7%/ $E^{1/4}$	1983
Bi <sub>4</sub> Ge <sub>3</sub> O <sub>12</sub> (BGO) (L3)	22 $X_0$	2%/ $\sqrt{E} \oplus 0.7\%$	1993
CsI (KTeV)	27 $X_0$	2%/ $\sqrt{E} \oplus 0.45\%$	1996
CsI(Tl) (BaBar)	16–18 $X_0$	2.3%/ $E^{1/4} \oplus 1.4\%$	1999
CsI(Tl) (BELLE)	16 $X_0$	1.7% for $E_\gamma > 3.5$ GeV	1998
PbWO <sub>4</sub> (PWO) (CMS)	25 $X_0$	3%/ $\sqrt{E} \oplus 0.5\% \oplus 0.2/E$	1997
Lead glass (OPAL)	20.5 $X_0$	5%/ $\sqrt{E}$	1990
Liquid Kr (NA48)	27 $X_0$	3.2%/ $\sqrt{E} \oplus 0.42\% \oplus 0.09/E$	1998
Scintillator/depleted U (ZEUS)	20–30 $X_0$	18%/ $\sqrt{E}$	1988
Scintillator/Pb (CDF)	18 $X_0$	13.5%/ $\sqrt{E}$	1988
Scintillator fiber/Pb spaghetti (KLOE)	15 $X_0$	5.7%/ $\sqrt{E} \oplus 0.6\%$	1995
Liquid Ar/Pb (NA31)	27 $X_0$	7.5%/ $\sqrt{E} \oplus 0.5\% \oplus 0.1/E$	1988
Liquid Ar/Pb (SLD)	21 $X_0$	8%/ $\sqrt{E}$	1993
Liquid Ar/Pb (H1)	20–30 $X_0$	12%/ $\sqrt{E} \oplus 1\%$	1998
Liquid Ar/depl. U (DØ)	20.5 $X_0$	16%/ $\sqrt{E} \oplus 0.3\% \oplus 0.3/E$	1993
Liquid Ar/Pb accordion (ATLAS)	25 $X_0$	10%/ $\sqrt{E} \oplus 0.4\% \oplus 0.3/E$	1996

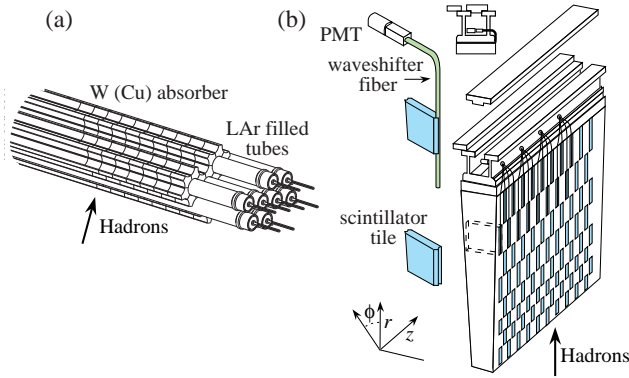
### 34.9.2. Hadronic calorimeters : [1–5,153]

Revised September 2013 by D. E. Groom (LBNL).

Hadronic calorimetry is considerably more difficult than EM calorimetry. For the same cascade containment fraction discussed in the previous section, the calorimeter would need to be  $\sim 30$  times deeper. Electromagnetic energy deposit from the decay of a small number of  $\pi^0$ 's are usually detected with greater efficiency than are the hadronic parts of the cascade, themselves subject to large fluctuations in neutron production, undetectable energy loss to nuclear disassociation, and other effects.

Most large hadron calorimeters are parts of large  $4\pi$  detectors at colliding beam facilities. At present these are sampling calorimeters: plates of absorber (Fe, Pb, Cu, or occasionally U or W) alternating with plastic scintillators (plates, tiles, bars), liquid argon (LAr), or gaseous detectors. The ionization is measured directly, as in LAr calorimeters, or via scintillation light observed by photodetectors (usually PMT's or silicon photodiodes). Wavelength-shifting fibers are often used to solve difficult problems of geometry and light collection uniformity. Silicon sensors are being studied for ILC detectors; in this case  $e-h$  pairs are collected. There are as many variants of these schemes as there are calorimeters, including variations in geometry of the absorber and sensors, *e.g.*, scintillating fibers threading an absorber [156], and the “accordion” LAr detector [157]. The latter has zig-zag absorber plates to minimize channeling effects; the

calorimeter is hermetic (no cracks), and plates are oriented so that cascades cross the same plate repeatedly. Another departure from the traditional sandwich structure is the LAr-tube design shown in Fig. 34.22(a) [158].



**Figure 34.22:** (a) ATLAS forward hadronic calorimeter structure (FCal2, 3) [158]. Tubes containing LAr are embedded in a mainly tungsten matrix. (b) ATLAS central calorimeter wedge; iron with plastic scintillator tile with wavelength-shifting fiber readout [159].

A relatively new variant in hadron calorimetry is the detection of Cerenkov light. Such a calorimeter is sensitive to relativistic  $e^\pm$ s in the EM showers plus a few relativistic pions. An example is the radiation-hard forward calorimeter in CMS, with iron absorber and quartz fiber readout by PMT's [160].

Ideally the calorimeter is segmented in  $\phi$  and  $\theta$  (or  $\eta = -\ln \tan(\theta/2)$ ). Fine segmentation, while desirable, is limited by cost, readout complexity, practical geometry, and the transverse size of the cascades—but see Ref. 161. An example, a wedge of the ATLAS central barrel calorimeter, is shown in Fig. 34.22(b) [159].

Much of the following discussion assumes an idealized calorimeter, with the same structure throughout and without leakage. “Real” calorimeters usually have an EM detector in front and a coarse “catcher” in the back. Complete containment is generally impractical.

In an inelastic hadronic collision a significant fraction  $f_{em}$  of the energy is removed from further hadronic interaction by the production of secondary  $\pi^0$ s and  $\eta$ s, whose decay photons generate high-energy electromagnetic (EM) showers. Charged secondaries ( $\pi^\pm$ ,  $p$ , ...) deposit energy via ionization and excitation, but also interact with nuclei, producing spallation protons and neutrons, evaporation neutrons, and spallation products. The charged collision products produce detectable ionization, as do the showering  $\gamma$ -rays from the prompt de-excitation of highly excited nuclei. The recoiling nuclei generate little or no detectable signal. The neutrons lose kinetic energy in elastic collisions, thermalize on a time scale of several  $\mu$ s, and are captured, with the production of more  $\gamma$ -rays—usually outside the acceptance gate of the electronics. Between endothermic spallation losses, nuclear recoils, and late neutron capture, a significant fraction of the hadronic energy (20%–40%, depending on the absorber and energy of the incident particle) is used to overcome nuclear binding energies and is therefore lost or “invisible.”

In contrast to EM showers, hadronic cascade processes are characterized by the production of relatively few high-energy particles. The lost energy and  $f_{em}$  are highly variable from event to event. Until there is event-by-event knowledge of both the EM fraction and the invisible energy loss, the energy resolution of a hadron calorimeter will remain significantly worse than that of its EM counterpart.

The efficiency  $e$  with which EM deposit is detected varies from event to event, but because of the large multiplicity in EM showers the variation is small. In contrast, because a variable fraction of the hadronic energy deposit is detectable, the efficiency  $h$  with which hadronic energy is detected is subject to considerably larger fluctuations. It thus makes sense to consider the ratio  $h/e$  as a stochastic variable.

Most energy deposit is by very low-energy electrons and charged hadrons. Because so many generations are involved in a high-energy cascade, the hadron spectra in a given material are essentially independent of energy except for overall normalization [163]. For this reason  $\langle h/e \rangle$  is a robust concept, independently of hadron energy and species.

If the detection efficiency for the EM sector is  $e$  and that for the hadronic sector is  $h$ , then the ratio of the mean response to a pion relative to that for an electron is

$$\langle \pi/e \rangle = \langle f_{em} \rangle + \langle f_h \rangle \langle h/e \rangle^* = 1 - (1 - \langle h/e \rangle) \langle f_h \rangle \quad (34.34)$$

It has been shown by a simple induction argument and verified by experiment, that the decrease in the average value of the hadronic energy fraction  $\langle f_h \rangle = 1 - \langle f_{em} \rangle$  as the projectile energy  $E$  increases is fairly well described by the power law [162,163]

$$\langle f_h \rangle \approx (E/E_0)^{m-1} \quad (\text{for } E > E_0), \quad (34.35)$$

at least up to a few hundred GeV. The exponent  $m$  depends logarithmically on the mean multiplicity and the mean fractional loss to  $\pi^0$  production in a single interaction. It is in the range 0.80–0.87.  $E_0$ , roughly the energy for the onset of inelastic collisions, is 1 GeV or a little less for incident pions [162]. Both  $m$  and  $E_0$  must be obtained experimentally for a given calorimeter configuration.

Only the product  $(1 - \langle h/e \rangle) E_0^{1-m}$  can be obtained by measuring  $\langle \pi/e \rangle$  as a function of energy. Since  $1 - m$  is small and  $E_0 \approx 1$  GeV for pion-induced cascades, this fact is usually ignored and  $\langle h/e \rangle$  is reported.

In a hadron-nucleus collision a large fraction of the incident energy is carried by a “leading particle” with the same quark content as the incident hadron. If the projectile is a charged pion, the leading particle is usually a pion, which can be neutral and hence contributes to the EM sector. This is not true for incident protons. The result is an increased mean hadronic fraction for incident protons:  $E_0 \approx 2.6$  GeV [162–165].

By definition,  $0 \leq f_{em} \leq 1$ . Its variance  $\sigma_{f_{em}}^2$  changes only slowly with energy, but perforce  $\langle f_{em} \rangle \rightarrow 1$  as the projectile energy increases. An empirical power law (unrelated to Eq. (34.34)) of the form  $\sigma_{f_{em}} = (E/E_1)^{1-\ell}$  (where  $\ell < 1$ ) describes the energy dependence of the variance adequately and has the right asymptotic properties [153]. For  $\langle h/e \rangle \neq 1$  (noncompensation), fluctuations in  $f_{em}$  significantly contribute to or even dominate the resolution. Since the  $f_{em}$  distribution has a high-energy tail, the calorimeter response is non-Gaussian with a high-energy tail if  $\langle h/e \rangle < 1$ . Noncompensation thus seriously degrades resolution and produces a nonlinear response.

It is clearly desirable to *compensate* the response, i.e., to design the calorimeter such that  $\langle h/e \rangle = 1$ . This is possible only with a sampling calorimeter, where several variables can be chosen or tuned:

1. Decrease the EM sensitivity. EM cross sections increase with  $Z$ ,<sup>†</sup> and most of the energy in an EM shower is deposited by low-energy electrons. A disproportionate fraction of the EM energy is thus deposited in the higher- $Z$  absorber. Lower- $Z$  cladding, such as the steel cladding on ZEUS U plates, preferentially absorbs low-energy  $\gamma$ s in EM showers and thus also lowers the electronic response. G10 signal boards in the DØ calorimeters and G10 next to silicon readout detectors has the same effect. The degree of EM signal suppression can be somewhat controlled by tuning the sensor/absorber thickness ratio.
2. Increase the hadronic sensitivity. The abundant neutrons produced in the cascade have large  $n$ - $p$  elastic scattering cross sections, so that low-energy scattered protons are produced in hydrogenous sampling materials such as butane-filled proportional counters or plastic scintillator. (The maximal fractional energy loss when a neutron scatters from a nucleus with mass number  $A$  is

\* Technically, we should write  $\langle f_h(h/e) \rangle$ , but we approximate it as  $\langle f_h \rangle \langle h/e \rangle$  to facilitate the rest of the discussion.

<sup>†</sup> The asymptotic pair-production cross section scales roughly as  $Z^{0.75}$ , and  $|dE/dx|$  slowly decreases with increasing  $Z$ .



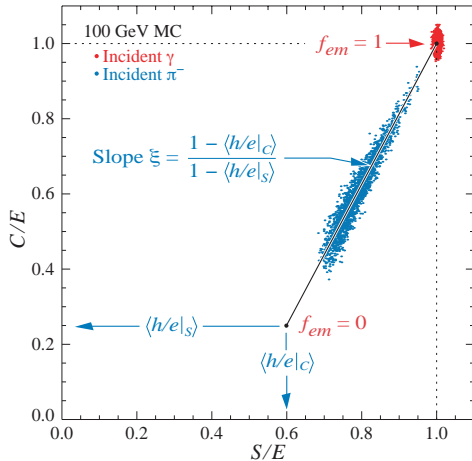
$4A/(1+A)^2$ .) The down side in the scintillator case is that the signal from a highly-ionizing stopping proton can be reduced by as much as 90% by recombination and quenching parameterized by Birks' Law (Eq. (34.2)).

3. Fabjan and Willis proposed that the additional signal generated in the aftermath of fission in  $^{238}\text{U}$  absorber plates should compensate nuclear fluctuations [166]. The production of fission fragments due to fast  $n$  capture was later observed [167]. However, while a very large amount of energy is released, it is mostly carried by low-velocity, very highly ionizing fission fragments which produce very little observable signal because of recombination and quenching. But in fact much of the compensation observed with the ZEUS  $^{238}\text{U}$ /scintillator calorimeter was mainly the result of methods 1 and 2 above.

Motivated very much by the work of Brau, Gabriel, Brückmann, and Wigmans [168], several groups built calorimeters which were very nearly compensating. The degree of compensation was sensitive to the acceptance gate width, and so could be somewhat further tuned. These included

- a) HELIOS with 2.5 mm thick scintillator plates sandwiched between 2 mm thick  $^{238}\text{U}$  plates (one of several structures);  $\sigma/E = 0.34/\sqrt{E}$  was obtained,
- b) ZEUS, 2.6 cm thick scintillator plates between 3.3 mm  $^{238}\text{U}$  plates;  $\sigma/E = 0.35/\sqrt{E}$ ,
- c) a ZEUS prototype with 10 mm Pb plates and 2.5 mm scintillator sheets;  $\sigma/E = 0.44/\sqrt{E}$ , and
- d) DØ, where the sandwich cell consists of a 4–6 mm thick  $^{238}\text{U}$  plate, 2.3 mm LAr, a G-10 signal board, and another 2.3 mm LAr gap;  $\sigma/E \approx 0.45/\sqrt{E}$ .

Given geometrical and cost constraints, the calorimeters used in modern collider detectors are not compensating:  $\langle h/e \rangle \approx 0.7$ , for the ATLAS central barrel calorimeter, is typical.



**Figure 34.23:** Dotplot of Monte Carlo  $C$  (Cherenkov) vs  $S$  (scintillator) signals for individual events in a dual readout calorimeter. Hadronic ( $\pi^-$ ) induced events are shown in blue, and scatter about the indicated event locus. Electromagnetic events cluster about  $(C,S) = (0,0)$ . In this case worse resolution (fewer p.e.'s) was assumed for the Cherenkov events, leading to the “elliptical” distribution.

A more versatile approach to compensation is provided by a *dual-readout calorimeter*, in which the signal is sensed by two readout systems with highly contrasting  $\langle h/e \rangle$ . Although the concept is more than two decades old [169], it was only recently been implemented by the DREAM collaboration [170]. The test beam calorimeter consisted of copper tubes, each filled with scintillator and quartz fibers. If the two signals  $C$  and  $S$  (quartz and scintillator) are both normalized to electron response, then for each event Eq. (34.34) takes the form

$$\begin{aligned} C &= E[f_{em} + \langle h/e \rangle_C(1 - f_{em})] \\ S &= E[f_{em} + \langle h/e \rangle_S(1 - f_{em})] \end{aligned} \quad (34.36)$$

for the Cherenkov and scintillator responses. On a dotplot of  $C/E$  vs  $S/E$ , events scatter about a line-segment locus described in Fig. 34.23. With increasing energy the distribution moves upward along the locus and becomes tighter. Equations 34.36 are linear in  $1/E$  and  $f_{em}$ , and are easily solved to obtain estimators of the *corrected* energy and  $f_{em}$  for each event. Both are subject to resolution effects, but contributions due to fluctuations in  $f_{em}$  are eliminated. The solution for the corrected energy is given by [163]:

$$E = \frac{\xi S - C}{\xi - 1}, \text{ where } \xi = \frac{1 - \langle h/e \rangle_C}{1 - \langle h/e \rangle_S} \quad (34.37)$$

$\xi$  is the energy-independent slope of the event locus on a plot of  $C$  vs  $S$ . It can be found either from the fitted slope or by measuring  $\pi/e$  as a function of  $E$ . Because we have no knowledge of  $h/e$  on an event-by-event basis, it has been replaced by  $\langle h/e \rangle$  in Eq. (34.37).  $\xi$  must be as far from unity as possible to optimize resolution, which means in practical terms that the scintillator readout of the calorimeter must be as compensating as possible.

Although the usually-dominant contribution of the  $f_{em}$  distribution to the resolution can be minimized by compensation or the use of dual calorimetry, there remain significant contributions to the resolution:

1. Incomplete corrections for leakage, differences in light collection efficiency, and electronics calibration.
2. Readout transducer shot noise (usually photoelectron statistics), plus electronic noise.
3. Sampling fluctuations. Only a small part of the energy deposit takes place in the scintillator or other sensor, and that fraction is subject to large fluctuations. This can be as high as  $40\%/\sqrt{E}$  (lead/scintillator). It is even greater in the Fe/scint case because of the very small sampling fraction (if the calorimeter is to be compensating), and substantially lower in a U/scint calorimeter. It is obviously zero for a homogeneous calorimeter.
4. Intrinsic fluctuations. The many ways ionization can be produced in a hadronic shower have different detection efficiencies and are subject to stochastic fluctuations. In particular, a very large fraction of the hadronic energy ( $\sim 20\%$  for Fe/scint,  $\sim 40\%$  for U/scint) is “invisible,” going into nuclear dissociation, thermalized neutrons, etc. The lost fraction depends on readout—it will be greater for a Cherenkov readout, less for an organic scintillator readout.

Except in a sampling calorimeter especially designed for the purpose, sampling and intrinsic resolution contributions cannot be separated. This may have been best studied by Drews *et al.* [171], who used a calorimeter in which even- and odd-numbered scintillators were separately read out. Sums and differences of the variances were used to separate sampling and intrinsic contributions.

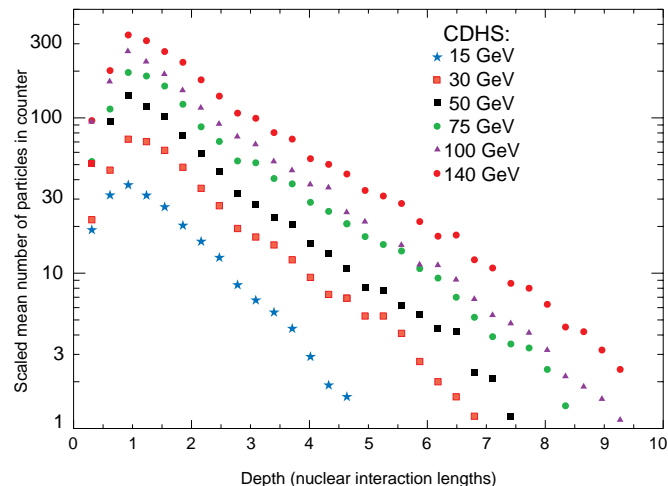
The fractional energy resolution can be represented by

$$\frac{\sigma}{E} = \frac{a_1(E)}{\sqrt{E}} \oplus \left| 1 - \left\langle \frac{h}{e} \right\rangle \right| \left( \frac{E}{E_1} \right)^{1-\ell} \quad (34.38)$$

The coefficient  $a_1$  is expected to have mild energy dependence for a number of reasons. For example, the sampling variance is  $(\pi/e)E$  rather than  $E$ . The term  $(E/E_1)^{1-\ell}$  is the parametrization of  $\sigma_{f_{em}}$  discussed above. Usually a plot of  $(\sigma/E)^2$  vs  $1/E$  is well-described by a straight line (constant  $a_1$ ) with a finite intercept—the square of the right term in Eq. (34.38), is called “the constant term.” Precise data show the slight downturn [156].

After the first interaction of the incident hadron, the average longitudinal distribution rises to a smooth peak. The peak position increases slowly with energy. The distribution becomes nearly exponential after several interaction lengths. Examples from the CDHS magnetized iron-scintillator sandwich calorimeter test beam calibration runs [172] are shown in Fig. 34.24. Proton-induced cascades are somewhat shorter and broader than pion-induced cascades [165]. A gamma distribution fairly well describes the longitudinal development of an EM shower, as discussed in Sec. 33.5. Following this logic, Bock *et al.* suggested that the profile of a hadronic cascade could be fitted by the sum of two  $\Gamma$  distributions, one with

a characteristic length  $X_0$  and the other with length  $\lambda_I$  [173]. Fits to this 4-parameter function are commonly used, *e.g.*, by the ATLAS Tilecal collaboration [165]. If the interaction point is not known (the usual case), the distribution must be convoluted with an exponential in the interaction length of the incident particle. Adragna *et al.* give an analytic form for the convoluted function [165].



**Figure 34.24:** Mean profiles of  $\pi^+$  (mostly) induced cascades in the CDHS neutrino detector [172]. Corresponding results for the ATLAS tile calorimeter can be found in Ref. 165.

The transverse energy deposit is characterized by a central core dominated by EM cascades, together with a wide “skirt” produced by wide-angle hadronic interactions [174].

The CALICE collaboration has tested a “tracking” calorimeter (AHCAL) with highly granular scintillator readout [161]. Since the position of the first interaction is observed, the average longitudinal and radial shower distributions are obtained.

While the average distributions might be useful in designing a calorimeter, they have little meaning for individual events, whose distributions are extremely variable because of the small number of particles involved early in the cascade.

Particle identification, primarily  $e\pi$  discrimination, is accomplished in most calorimeters by depth development. An EM shower is mostly contained in  $15X_0$  while a hadronic shower takes about  $4\lambda_I$ . In high- $A$  absorbers such as Pb,  $X_0/\lambda_I \sim 0.03$ . In a fiber calorimeter, such as the RD52 dual-readout calorimeter [175],  $e\pi$  discrimination is achieved by differences in the Cerenkov and scintillation signals, lateral spread, and timing differences, ultimately achieving about 500:1 discrimination.

### 34.9.3. Free electron drift velocities in liquid ionization chambers :

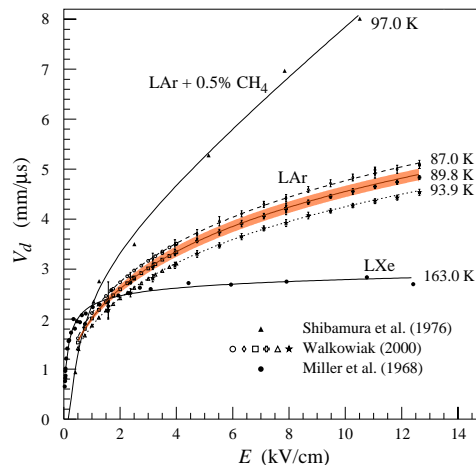
Written August 2009 by W. Walkowiak (U. Siegen)

Drift velocities of free electrons in LAr [176] are given as a function of electric field strength for different temperatures of the medium in Fig. 34.25. The drift velocities in LAr have been measured using a double-gridded drift chamber with electrons produced by a laser pulse on a gold-plated cathode. The average temperature gradient of the drift velocity of the free electrons in LAr is described [176] by

$$\frac{\Delta v_d}{\Delta T v_d} = (-1.72 \pm 0.08) \% / \text{K}.$$

Earlier measurements [177–180] used different techniques and show systematic deviations of the drift velocities for free electrons which cannot be explained by the temperature dependence mentioned above.

Drift velocities of free electrons in LXe [178] as a function of electric field strength are also displayed in Fig. 34.25. The drift velocity saturates for  $|\mathbf{E}| > 3 \text{ kV/cm}$ , and decreases with increasing temperature for LXe as well as measured *e.g.* by [181].



**Figure 34.25:** Drift velocity of free electrons as a function of electric field strength for LAr [176], LAr + 0.5% CH<sub>4</sub> [178] and LXe [177]. The average temperatures of the liquids are indicated. Results of a fit to an empirical function [182] are superimposed. In case of LAr at 91 K the error band for the global fit [176] including statistical and systematic errors as well as correlations of the data points is given. Only statistical errors are shown for the individual LAr data points.

The addition of small concentrations of other molecules like N<sub>2</sub>, H<sub>2</sub> and CH<sub>4</sub> in solution to the liquid typically increases the drift velocities of free electrons above the saturation value [178,179], see example for CH<sub>4</sub> admixture to LAr in Fig. 34.25. Therefore, actual drift velocities are critically dependent on even small additions or contaminations.

## 34.10. Accelerator Neutrino Detectors

Written August 2015 by M.O. Wascko (Imperial College London).

### 34.10.1. Introduction :

Accelerator neutrino experiments span many orders of magnitude in neutrino energy, from a few MeV to hundreds of GeV. This wide range of neutrino energy is driven by the many physics applications of accelerator neutrino beams. Foremost among them is neutrino oscillation, which varies as the ratio  $L/E_\nu$ , where  $L$  is the neutrino baseline (distance travelled), and  $E_\nu$  is the neutrino energy. But accelerator neutrino beams have also been used to study the nature of the weak interaction, to probe nucleon form factors and structure functions, and to study nuclear structure.

The first accelerator neutrino experiment used neutrinos from the decays of high energy pions in flight to show that the neutrinos emitted from pion decay are different from the neutrinos emitted by beta decay [183]. The field of accelerator neutrino experiments did not expand beyond this until Simon van der Meer’s invention of the magnetic focusing horn [184], which significantly increased the flux of neutrinos aimed toward the detector. In this mini-review, we focus on experiments employing decay-in-flight beams—pions, kaons, charmed mesons, and taus—producing fluxes of neutrinos and antineutrinos from  $\sim 10 \text{ MeV}$  to  $\sim 100 \text{ GeV}$ .

Neutrino interactions with matter proceed only through the weak interaction, making the cross section extremely small and requiring high fluxes of neutrinos and large detector masses in order to achieve satisfactory event rates. Therefore, neutrino detector design is a balancing act taking into account sufficient numbers of nuclear targets (often achieved with inactive detector materials), adequate sampling/segmentation to ensure accurate reconstruction of the tracks and showers produced by neutrino-interaction secondary particles, and practical readout systems to allow timely analysis of data.

### 34.10.2. Signals and Backgrounds :

The neutrino interaction processes available increase with increasing neutrino energy as interaction thresholds are crossed; in general neutrino-interaction cross sections grow with energy; for a detailed discussion of neutrino interactions see [185]. The multiplicity of secondary particles from each interaction process grows in complexity with neutrino energy, while the forward-boost due to increasing  $E_\nu$  compresses the occupied phase space in the lab frame, impacting detector designs. Because decay-in-flight beams produce neutrinos at well-defined times, leading to very small duty factors, the predominant backgrounds stem from unwanted beam-induced neutrino interactions, i.e. neutrinos interacting via other processes than the one being studied. This becomes increasingly true at high energies because the secondary particles produced by neutrino interactions yield detector signals that resemble cosmic backgrounds less and less.

Below, we describe a few of the dominant neutrino interaction processes, with a focus on the final state particle content and topologies.

#### 34.10.2.1. Charged-Current Quasi-Elastic Scattering and Pions:

Below  $\sim 2$  GeV neutrino energy, the dominant neutrino-nucleus interaction process is quasi-elastic (QE) scattering. In the charged current (CC) mode, the CCQE base neutrino reaction is  $\nu_\ell n \rightarrow \ell^- p$ , where  $\ell = e, \mu, \tau$ , and similarly for antineutrinos,  $\bar{\nu}_\ell p \rightarrow \ell^+ n$ . The final state particles are a charged lepton, and perhaps a recoiling nucleon if it is given enough energy to escape the nucleus. Detectors designed to observe this process should have good single-particle track resolution for muon neutrino interactions, but should have good  $\mu/e$  separation for electron neutrino interactions. Because the interaction cross section falls sharply with  $Q^2$ , the lepton typically carries away more of the neutrino's kinetic energy than the recoiling nucleon. The fraction of backward-scattered leptons is large, however, so detectors with  $4\pi$  coverage are desirable. The dominant backgrounds in this channel tend to come from single pion production events in which the pion is not detected.

Near 1 GeV, the quasi-elastic cross section is eclipsed by pion production processes. A typical single pion production (CC $1\pi$ ) reaction is  $\nu_\ell n \rightarrow \ell^- \pi^+ n$ , but many more final state particle combinations are possible. Single pion production proceeds through the coherent channel and many incoherent processes, dominated by resonance production. With increasing neutrino energy, higher-order resonances can be excited, leading to multiple pions in the final state. Separating these processes from quasi-elastic scattering, and indeed from each other, requires tagging, and ideally reconstructing, the pions. Since these processes can produce neutral pions, electromagnetic (EM) shower reconstruction is more important here than it is for the quasi-elastic channel. The predominant backgrounds for pion production change with increasing neutrino energy. Detection of pion processes is also complicated because near threshold the quasi-elastic channel creates pion backgrounds through final state interactions of the recoiling nucleon, and at higher energies backgrounds come from migration of multiple pion events in which one or more pions is not detected.

#### 34.10.2.2. Deep Inelastic Scattering:

Beyond a few GeV, the neutrino has enough energy to probe the nucleon at the parton scale, leading to deep inelastic scattering (DIS). In the charged-current channel, the DIS neutrino reaction is  $\nu_\ell N \rightarrow \ell^- X$ , where  $N$  is a nucleon and  $X$  encompasses the entire recoiling hadronic system. The final state particle reconstruction revolves around accurate reconstruction of the lepton momentum and containment and reconstruction of the hadronic shower energy. Because of the high neutrino energies involved, DIS events are very forward boosted, and can have extremely long particle tracks. For this reason, detectors measuring DIS interactions must be large to contain the hadronic showers in the detector volume.

#### 34.10.2.3. Neutral Currents:

Neutrino interactions proceeding through the neutral current (NC) channel are identified by the lack of a charged lepton in the final state. For example, the NC elastic reaction is  $\nu_\ell N \rightarrow \nu_\ell N$ , and the NC DIS reaction is  $\nu_\ell N \rightarrow \nu_\ell X$ . NC interactions are suppressed relative to CC interactions by a factor involving the weak mixing angle; the primary backgrounds for NC interactions come from CC interactions in which the charged lepton is misidentified.

#### 34.10.3. Instances of Neutrino Detector Technology :

Below we describe many of the actual detectors that have been built and operated for use in accelerator neutrino beams.

##### 34.10.3.1. Spark Chambers:

In the first accelerator neutrino beam experiment, Lederman, Schwartz, and Steinberger [183] used an internally-triggered spark chamber detector, filled with 10 tons of Al planes and surrounded by external scintillator veto planes, to distinguish muon tracks from electron showers, and hence muon neutrinos from electron neutrinos. The inactive Al planes served as the neutrino interaction target and as radiators for EM shower development. The detector successfully showed the presence of muon tracks from neutrino interactions. It was also sensitive to the hadronic showers induced by NC interactions, which were unknown at the time. More than a decade later, the Aachen-Padova [186] experiment at CERN also employed an Al spark chamber to detect  $\sim 2$  GeV neutrinos.

##### 34.10.3.2. Bubble Chambers:

Several large bubble chamber detectors were employed as accelerator neutrino detectors in the 1970s and 80s, performing many of the first studies of the properties of the weak interaction. Bubble chambers provide exquisite granularity in the reconstruction of secondary particles, allowing very accurate separation of interaction processes. However, the extremely slow and labor-intensive acquisition and analysis of the data from photographic film led to them being phased out in favor of electronically read out detectors.

The Gargamelle [187] detector at CERN used Freon and propane gas targets to make the first observation of neutrino-induced NC interactions and more. The BEBC [188] detector at CERN was a bubble chamber that was alternately filled with liquid hydrogen, deuterium, and a neon-hydrogen mixture; BEBC was also outfitted with a track-sensitive detector to improve event tagging, and sometimes used with a small emulsion chamber. The SKAT [192] heavy freon bubble chamber was exposed to wideband neutrino and antineutrino beams at the Serpukhov laboratory in the former Soviet Union. A series of American bubble chambers in the 1970's and 1980's made measurements on free nucleons that are still crucial inputs for neutrino-nucleus scattering predictions. The 12-foot bubble chamber at ANL [189] in the USA used both deuterium and hydrogen targets, as did the 7-foot bubble chamber at BNL [190]. Fermilab's 15 foot bubble chamber [191] used deuterium and heavy neon targets.

##### 34.10.3.3. Iron Tracking Calorimeters:

Because of the forward boost of high energy interactions, long detectors made of magnetized iron interspersed with active detector layers have been very successfully employed. The long magnetized detectors allow measurements of the momentum of penetrating muons. The iron planes also act as shower-inducing layers, allowing separation of EM and hadronic showers; the large number of iron planes provide enough mass for high statistics and/or shower containment. Magnetized iron spectrometers have been used for studies of the weak interaction, measurements of structure functions, and searches for neutrino oscillation. Non-magnetized iron detectors have also been successfully employed as neutrino monitors for oscillation experiments and also for neutrino-nucleus interaction studies.

The CDHS [201] detector used layers of magnetized iron modules interspersed with wire drift chambers, with a total (fiducial) mass of 1250 t (750 t), to detect neutrinos in the range 30–300 GeV. Within each iron module, 5 cm (or 15 cm) iron plates were interspersed with scintillation counters. The FNAL Lab-E neutrino detector was used by the CCFR [202] and NuTeV [203] collaborations to

perform a series of experiments in the Fermilab high energy neutrino beam ( $50 \text{ GeV} < E_\nu < 300 \text{ GeV}$ ). The detector was comprised of six iron target calorimeter modules, with 690 t total target mass, followed by three muon spectrometer modules, followed by two drift chambers. Each iron target calorimeter module comprised 5.2 cm thick steel plates interspersed with liquid scintillation counters and drift chambers. The muon spectrometer was comprised of toroidal iron magnets interleaved with drift chambers. The MINOS [204] detectors, a near detector of 980 t at FNAL and a far detector of 5400 t in the Soudan mine, are functionally identical magnetized iron calorimeters, comprised of iron plates interleaved with layers of 4 cm wide plastic scintillator strips in alternating orientations. The T2K [222] on-axis detector, INGRID, consists of 16 non-magnetized iron scintillator sandwich detectors, each with nine 6.5 cm iron plane (7.1 t total) interspersed between layers of 5 cm wide plastic scintillator strips readout out by multi-pixel photon counters (MPPCs) coupled to WLS fibers. Fourteen of the INGRID modules are arranged in a cross-hair configuration centered on the neutrino beam axis.

#### 34.10.3.4. Cherenkov Detectors:

Open volume water Cherenkov detectors were originally built to search for proton decay. Large volumes of ultra-pure water were lined with photomultipliers to collect Cherenkov light emitted by the passage of relativistic charged particles. See Sec. 35.3.1 for a detailed discussion of deep liquid detectors for rare processes.

When used to detect  $\sim \text{GeV}$  neutrinos, the detector medium acts as a natural filter for final state particles below the Cherenkov threshold; this feature has been exploited successfully by the K2K, MiniBooNE (using mineral oil instead of water), and T2K neutrino oscillation experiments. However, at higher energies Cherenkov detectors become less accurate because the overlapping rings from many final state particles become increasingly difficult to resolve.

The second-generation Cherenkov detector in Japan, Super-Kamiokande [193] (Super-K), comprises 50 kt (22.5 kt fiducial) of water viewed by 11,146 50 cm photomultiplier tubes, giving 40% photocathode coverage; it is surrounded by an outer detector region viewed by 1,885 20 cm photomultipliers. Super-K is the far detector for K2K and T2K, and is described in greater detail elsewhere in this review. The K2K experiment also employed a 1 kt water Cherenkov detector in the suite of near detectors [194], with 690 photomultipliers (40% photocathode coverage) viewing the detector volume. The MiniBooNE detector at FNAL was a 0.8 kt [195] mineral oil Cherenkov detector, with 1,520 20 cm photomultipliers (10% photocathode coverage) surrounded by a veto detector with 240 20 cm photomultipliers.

#### 34.10.3.5. Scintillation Detectors:

Liquid and solid scintillator detectors also employ fully (or nearly fully) active detector media. Typically organic scintillators, which emit into the ultraviolet range, are dissolved in mineral oil or plastic and read out by photomultipliers coupled to wavelength shifters (WLS). Open volume scintillation detectors lined with photomultipliers are conceptually similar to Cherenkov detectors, although energy reconstruction is calorimetric in nature as opposed to kinematic (see also Sec. 35.3.1). For higher energies and higher particle multiplicities, it becomes beneficial to use segmented detectors to help distinguish particle tracks and showers from each other.

The LSND [197] detector at LANL was an open volume liquid scintillator detector (of mass 167 t) employed to detect relatively low energy ( $< 300 \text{ MeV}$ ) neutrinos. The NO $\nu$ A [200] detectors use segmented volumes of liquid scintillator in which the scintillation light is collected by WLS fibers in the segments that are coupled to avalanche photodiodes (APDs) at the ends of the volumes. The NO $\nu$ A far detector, located in Ash River, MN, is comprised of 896 layers of 15.6 m long extruded PVC scintillator cells for a total mass of 14 kt; the NO $\nu$ A near detector is comprised of 214 layers of 4.1 m scintillator volumes for a total mass of 300 t. Both are placed in the NuMI beamline at  $0.8^\circ$  off-axis. The SciBar (Scintillation Bar) detector was originally built for K2K at KEK in Japan and then re-used for SciBooNE [198] at FNAL. SciBar used plastic scintillator strips with  $1.5 \text{ cm} \times 2.5 \text{ cm}$  rectangular cross section, read out by multianode

photomultipliers (MAPMTs) coupled to WLS fibers, arranged in alternating horizontal and vertical layers, with a total mass of 15 t. Both SciBooNE and K2K employed an EM calorimeter downstream of SciBar and a muon range detector (MRD) downstream of that. The MINERvA [199] detector, in the NuMI beam at FNAL, utilizes a central tracker comprising 8.3 t of plastic scintillator strips with triangular cross section, and is also read out by MAPMTs coupled to WLS fibers. MINERvA employs several more subsystems and is described more fully below.

#### 34.10.3.6. Liquid Argon Time Projection Chambers:

Liquid argon time projection chambers (LAR-TPCs) were conceived in the 1970s as a way to achieve a fully active detector with sub-centimeter track reconstruction [205]. A massive volume of purified liquid argon is put under a strong electric field (hundreds of V/cm), so that the liberated electrons from the paths of ionizing particles can be drifted to the edge of the volume and read out, directly by collecting charge from wire planes or non-destructively through charge induction in the wire planes. A dual-phase readout method is also being developed, in which the charge is drifted vertically and then passed through an amplification region inside a gas volume above the liquid volume; the bottom of the liquid volume is equipped with a PMT array for detecting scintillation photons from the liquid argon. The first large scale LAR-TPC was the ICARUS T-600 module [206], comprising 760 t of liquid argon with a charge drift length of 1.5 m read out by wires with 3 mm pitch, which operated in LNGS, both standalone and also exposed to the CNGS high energy neutrino beam. The ArgoNeUT [207] detector at FNAL, with fiducial mass 25 kg of argon read out with 4 mm pitch wires, was exposed to the NuMI neutrino and antineutrino beams. The MicroBooNE [208] detector at FNAL comprises 170 t of liquid Ar, read out with 3 mm wire pitch, which began collecting data in the Booster Neutrino Beam Oct 2015. A LAR-TPC has also been chosen as the detector design for the future DUNE neutrino oscillation experiment, from FNAL to Sanford Underground Research Facility; both single and dual phase modules are planned.

#### 34.10.3.7. Emulsion Detectors:

Photographic film emulsions have been employed in particle physics experiments since the 1940s [209]. Thanks to advances in scanning technology and automation [213], they have been successfully employed as neutrino detectors. Emulsions are used for experiments observing CC tau neutrino interactions, where the short lifetime of the tau,  $\tau_\tau = 2.90 \times 10^{-13} \text{ s}$ , leading to the short mean path length,  $c \times \tau = 87 \mu\text{m}$ , requires extremely precise track resolution. They are employed in hybrid detectors in which the emulsion bricks are embedded inside fine-grained tracker detectors. In the data analysis, the tracker data are used to select events with characteristics typical of a tau decay in the final state, such as missing energy and unbalanced transverse momentum. The reconstructed tracks are projected back into an emulsion brick and used as the search seed for a neutrino interaction vertex.

E531 [210] at Fermilab tested many of the emulsion-tracker hybrid techniques employed by later neutrino experiments, in a detector with approximately 9 kg of emulsion target. The CHORUS [211] experiment at CERN used 1,600 kg of emulsion, in a hybrid detector with a fiber tracker, high resolution calorimeter, and muon spectrometer, to search for  $\nu_\mu \rightarrow \nu_\tau$  oscillation. The DONuT [212] experiment at FNAL used a hybrid detector, with 260 kg of emulsion bricks interspersed with fiber trackers, followed by a magnetic spectrometer, and calorimeter, to make the first direct observation of tau neutrino CC interactions. More recently, the OPERA [214,215,216] experiment used an automated hybrid emulsion detector, with 1,300 t of emulsion, to make the first direct observation of the appearance of  $\nu_\tau$  in a  $\nu_\mu$  beam.

### 34.10.3.8. Hybrid Detectors:

The CHARM detector [217] at CERN was built to study neutral-current interactions and search for muon neutrino oscillation. It was a fine-grained ionization calorimeter tracker with approximately 150 t of marble as neutrino target, surrounded by a magnetized iron muon system for tagging high angle muons, and followed downstream by a muon spectrometer. The CHARM II detector [218] at CERN comprised a target calorimeter followed by a downstream muon spectrometer. Each target calorimeter module consists of a 4.8 cm thick glass plate followed by a layer of plastic streamer tubes, with spacing 1 cm, instrumented with 2 cm wide pickup strips. Every fifth module is followed by a 3 cm thick scintillator layer. The total mass of the target calorimeter was 692 t.

The Brookhaven E-734 [219] detector was a tracking calorimeter made up of 172 t liquid scintillator modules interspersed with proportional drift tubes, followed by a dense EM calorimeter and a muon spectrometer downstream of that. The detector was exposed to a wideband horn-focused beam with peak neutrino energy near 1 GeV. The Brookhaven E-776 [220] experiment comprised a finely segmented EM calorimeter, with 2.54 cm concrete absorbers interspersed with planes of drift tubes and acrylic scintillation counters, with total mass 240 t, followed by a muon spectrometer.

The NOMAD [221] detector at CERN consisted of central tracker detector inside a 0.4 T dipole magnet (the magnet was originally used by the UA1 experiment at CERN) followed by a hadronic calorimeter and muon detectors downstream of the magnet. The main neutrino target is 3 t of drift chambers followed downstream by transition radiation detectors which are followed by an EM calorimeter. NOMAD was exposed to the same wideband neutrino beam as was CHORUS.

MINERvA, introduced above, is, in its entirety, a hybrid detector, based around a central plastic scintillator tracker. The scintillator tracker is surrounded by electromagnetic and hadronic calorimetry, which is achieved by interleaving thin lead (steel) layers between the scintillator layers for the ECAL (HCAL). MINERvA is situated upstream of the MINOS near detector which acts as a muon spectrometer. Upstream of the scintillator tracker is a nuclear target region containing inactive layers of C (graphite), Pb, Fe (steel), and O (water). MINERvA's physics goals span a wide range of neutrino-nucleus interaction studies, from form factors to nuclear effects.

T2K [222] in Japan employs two near detectors at 280 m from the neutrino beam target, one centered on the axis of the horn-focused J-PARC neutrino beam and one placed 2.5° off-axis. The on-axis detector, INGRID, is described above. The 2.5° off-axis detector, ND280, employs the UA1 magnet (at 0.2 T) previously used by NOMAD. Inside the magnet volume are three separate detector systems: the trackers, the Pi0 Detector (P0D), and several ECAL modules. The tracker detectors comprise two fine-grained scintillator detectors (FGDs), read out by MPPCs coupled to WLS fibers, interleaved between three gas TPCs read out by micromegas planes. The downstream FGD contains inactive water layers in addition to the scintillators. Upstream of the tracker is the P0D, a sampling tracker calorimeter with active detector materials comprising plastic scintillator read out by MPPCs and WLS fibers, and inactive sheets of brass radiators and refillable water modules. Surrounding the tracker and P0D, but still inside the magnet, are lead-scintillator EM sampling calorimeters.

### 34.11. Superconducting magnets for collider detectors

Revised September 2015 by Y. Makida (KEK)

**34.11.1. Solenoid Magnets :** In all cases SI unit are assumed, so that the magnetic field,  $B$ , is in Tesla, the stored energy,  $E$ , is in joules, the dimensions are in meters, and  $\mu_0 = 4\pi \times 10^{-7}$ .

The magnetic field ( $B$ ) in an ideal solenoid with a flux return iron yoke, in which the magnetic field is  $< 2$  T, is given by

$$B = \frac{\mu_0 n I}{L} \quad (34.39)$$

where  $n$  is the number of turns,  $I$  is the current and  $L$  is the coil length. In an air-core solenoid, the central field is given by

$$B(0,0) = \mu_0 n I \frac{L}{\sqrt{L^2 + 4R^2}}, \quad (34.40)$$

where  $R$  is the coil radius.

In most cases, momentum analysis is made by measuring the circular trajectory of the passing particles according to  $p = mv = q r B$ , where  $p$  is the momentum,  $m$  the mass,  $q$  the charge,  $r$  the bending radius. The sagitta,  $s$ , of the trajectory is given by

$$s = q B \ell^2 / 8p, \quad (34.41)$$

where  $\ell$  is the path length in the magnetic field. In a practical momentum measurement in colliding beam detectors, it is more effective to increase the magnetic volume than the field strength, since

$$dp/p \propto p/B \ell^2, \quad (34.42)$$

where  $\ell$  corresponds to the solenoid coil radius  $R$ . The energy stored in the magnetic field of any magnet is calculated by integrating  $B^2$  over all space:

$$E = \frac{1}{2\mu_0} \int B^2 dV \quad (34.43)$$

If the coil thin and inside an iron return yoke, (which is the case if it is to superconducting coil), then

$$E \approx (B^2/2\mu_0)\pi R^2 L. \quad (34.44)$$

For a detector in which the calorimetry is outside the aperture of the solenoid, the coil must be thin in terms of radiation and absorption lengths. This usually means that the coil is superconducting and that the vacuum vessel encasing it is of minimum real thickness and fabricated of a material with long radiation length. There are two major contributors to the thickness of a thin solenoid:

- 1) The conductor consisting of the current-carrying superconducting material (usually Nb-Ti/Cu) and the quench protecting stabilizer (usually aluminum) are wound on the inside of a structural support cylinder (usually aluminum also). The coil thickness scales as  $B^2 R$ , so the thickness in radiation lengths ( $X_0$ ) is

$$t_{\text{coil}}/X_0 = (R/\sigma_h X_0)(B^2/2\mu_0), \quad (34.45)$$

where  $t_{\text{coil}}$  is the physical thickness of the coil,  $X_0$  the average radiation length of the coil/stabilizer material, and  $\sigma_h$  is the hoop stress in the coil [225].  $B^2/2\mu_0$  is the magnetic pressure. In large detector solenoids, the aluminum stabilizer and support cylinders dominate the thickness; the superconductor (Nb-Ti/Cu) contributes a smaller fraction. The main coil and support cylinder components typically contribute about 2/3 of the total thickness in radiation lengths.

- 2) Another contribution to the material comes from the outer cylindrical shell of the vacuum vessel. Since this shell is susceptible to buckling collapse, its thickness is determined by the diameter, length and the modulus of the material of which it is fabricated. The outer vacuum shell represents about 1/3 of the total thickness in radiation length.

**Table 34.9:** Progress of superconducting magnets for particle physics detectors.

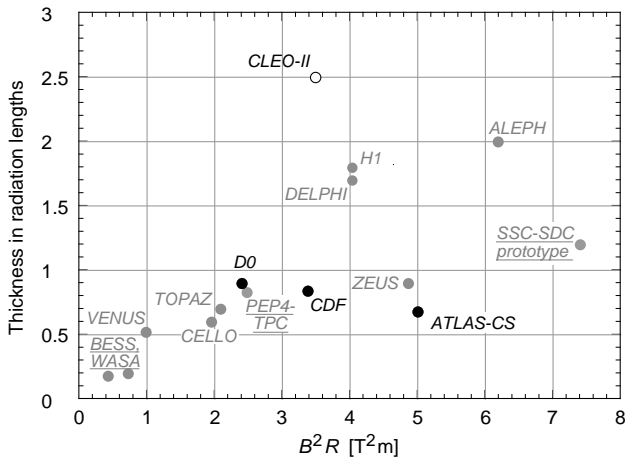
Experiment	Laboratory	$B$ [T]	Radius [m]	Length [m]	Energy [MJ]	$X/X_0$	$E/M$ [kJ/kg]
TOPAZ*	KEK	1.2	1.45	5.4	20	0.70	4.3
CDF*	Tsukuba/Fermi	1.5	1.5	5.07	30	0.84	5.4
VENUS*	KEK	0.75	1.75	5.64	12	0.52	2.8
AMY*	KEK	3	1.29	3	40	†	
CLEO-II*	Cornell	1.5	1.55	3.8	25	2.5	3.7
ALEPH*	Saclay/CERN	1.5	2.75	7.0	130	2.0	5.5
DELPHI*	RAL/CERN	1.2	2.8	7.4	109	1.7	4.2
ZEUS*	INFN/DESY	1.8	1.5	2.85	11	0.9	5.5
H1*	RAL/DESY	1.2	2.8	5.75	120	1.8	4.8
BaBar*	INFN/SLAC	1.5	1.5	3.46	27	†	3.6
D0*	Fermi	2.0	0.6	2.73	5.6	0.9	3.7
BELLE*	KEK	1.5	1.8	4	42	†	5.3
BES-III	IHEP	1.0	1.475	3.5	9.5	†	2.6
ATLAS-CS	ATLAS/CERN	2.0	1.25	5.3	38	0.66	7.0
ATLAS-BT	ATLAS/CERN	1	4.7–9.75	26	1080	(Toroid)†	
ATLAS-ET	ATLAS/CERN	1	0.825–5.35	5	$2 \times 250$	(Toroid)†	
CMS	CMS/CERN	4	6	12.5	2600	†	12
SiD**	ILC	5	2.9	5.6	1560	†	12
ILD**	ILC	4	3.8	7.5	2300	†	13
SiD**	CLIC	5	2.8	6.2	2300	†	14
ILD**	CLIC	4	3.8	7.9	2300	†	
FCC**		6	6	23	54000	†	12

\* No longer in service

\*\*Conceptual design in future

† EM calorimeter is inside solenoid, so small  $X/X_0$  is not a goal**34.11.2. Properties of collider detector magnets :**

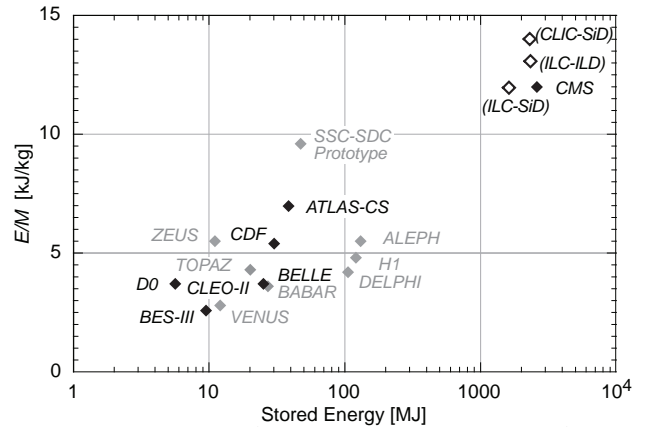
The physical dimensions, central field stored energy and thickness in radiation lengths normal to the beam line of the superconducting solenoids associated with the major collider are given in Table 34.9 [224]. Fig. 34.26 shows thickness in radiation lengths as a function of  $B^2R$  in various collider detector solenoids.



**Figure 34.26:** Magnet wall thickness in radiation length as a function of  $B^2R$  for various detector solenoids. Gray entries are for magnets no longer in use, and entries underlined are not listed in Table 34.9. Open circles are for magnets not designed to be “thin.” The SSC-SDC prototype provided important R&D for LHC magnets.

The ratio of stored energy to cold mass ( $E/M$ ) is a useful performance measure. It can also be expressed as the ratio of the stress,  $\sigma_h$ , to twice the equivalent density,  $\rho$ , in the coil [225]:

$$\frac{E}{M} = \frac{E}{\rho 2\pi t_{\text{coil}} RL} \approx \frac{\sigma_h}{2\rho} \quad (34.46)$$



**Figure 34.27:** Ratio of stored energy to cold mass for major detector solenoids. Gray indicates magnets no longer in operation.

The  $E/M$  ratio in the coil is approximately equivalent to  $H^*$ , the enthalpy of the coil, and it determines the average coil temperature rise after energy absorption in a quench:

$$E/M = H(T_2) - H(T_1) \approx H(T_2) \quad (34.47)$$

where  $T_2$  is the average coil temperature after the full energy absorption in a quench, and  $T_1$  is the initial temperature.  $E/M$  ratios of 5, 10, and 20 kJ/kg correspond to  $\sim 65$ ,  $\sim 80$ , and  $\sim 100$  K, respectively. The  $E/M$  ratios of various detector magnets are shown in Fig. 34.27 as a function of total stored energy. One would like the cold mass to be as small as possible to minimize the thickness,

\* The enthalpy, or heat content, is called  $H$  in the thermodynamics literature. It is not to be confused with the magnetic field intensity  $B/\mu$ .



but temperature rise during a quench must also be minimized. An  $E/M$  ratio as large as 12 kJ/kg is designed into the CMS solenoid, with the possibility that about half of the stored energy can go to an external dump resistor. Thus the coil temperature can be kept below 80 K if the energy extraction system works well. The limit is set by the maximum temperature that the coil design can tolerate during a quench. This maximum local temperature should be <130 K (50 K + 80 K), so that thermal expansion effects, which are remarkable beyond 80 K, in the coil are manageable less than 50 K.

### 34.11.3. Toroidal magnets :

Toroidal coils uniquely provide a closed magnetic field without the necessity of an iron flux-return yoke. Because no field exists at the collision point and along the beam line, there is, in principle, no effect on the beam. On the other hand, the field profile generally has  $1/r$  dependence. The particle momentum may be determined by measurements of the deflection angle combined with the sagitta. The deflection (bending) power  $BL$  is

$$BL \approx \int_{R_i}^{R_0} \frac{B_i R_i dR}{R \sin \theta} = \frac{B_i R_i}{\sin \theta} \ln(R_0/R_i), \quad (34.48)$$

where  $R_i$  is the inner coil radius,  $R_0$  is the outer coil radius, and  $\theta$  is the angle between the particle trajectory and the beam line axis. The momentum resolution given by the deflection may be expressed as

$$\frac{\Delta p}{p} \propto \frac{p}{BL} \approx \frac{p \sin \theta}{B_i R_i \ln(R_0/R_i)}. \quad (34.49)$$

The momentum resolution is better in the forward/backward (smaller  $\theta$ ) direction. The geometry has been found to be optimal when  $R_0/R_i \approx 3-4$ . In practical designs, the coil is divided into 6-12 lumped coils in order to have reasonable acceptance and accessibility. This causes the coil design to be much more complex. The mechanical structure needs to sustain the decentering force between adjacent coils, and the peak field in the coil is 3-5 times higher than the useful magnetic field for the momentum analysis [223].

### 34.12. Measurement of particle momenta in a uniform magnetic field [226,227]

The trajectory of a particle with momentum  $p$  (in GeV/c) and charge  $ze$  in a constant magnetic field  $\vec{B}$  is a helix, with radius of curvature  $R$  and pitch angle  $\lambda$ . The radius of curvature and momentum component perpendicular to  $\vec{B}$  are related by

$$p \cos \lambda = 0.3 z B R, \quad (34.50)$$

where  $B$  is in tesla and  $R$  is in meters.

The distribution of measurements of the curvature  $k \equiv 1/R$  is approximately Gaussian. The curvature error for a large number of uniformly spaced measurements on the trajectory of a charged particle in a uniform magnetic field can be approximated by

$$(\delta k)^2 = (\delta k_{\text{res}})^2 + (\delta k_{\text{ms}})^2, \quad (34.51)$$

where  $\delta k$  = curvature error

$\delta k_{\text{res}}$  = curvature error due to finite measurement resolution

$\delta k_{\text{ms}}$  = curvature error due to multiple scattering.

If many ( $\geq 10$ ) uniformly spaced position measurements are made along a trajectory in a uniform medium,

$$\delta k_{\text{res}} = \frac{\epsilon}{L'^2} \sqrt{\frac{720}{N+4}}, \quad (34.52)$$

where  $N$  = number of points measured along track

$L'$  = the projected length of the track onto the bending plane

$\epsilon$  = measurement error for each point, perpendicular to the trajectory.

If a vertex constraint is applied at the origin of the track, the coefficient under the radical becomes 320.

For arbitrary spacing of coordinates  $s_i$  measured along the projected trajectory and with variable measurement errors  $\epsilon_i$  the curvature error  $\delta k_{\text{res}}$  is calculated from:

$$(\delta k_{\text{res}})^2 = \frac{4}{w} \frac{V_{ss}}{V_{ss} V_{s^2 s^2} - (V_{ss^2})^2}, \quad (34.53)$$

where  $V$  are covariances defined as  $V_{s^m s^n} = \langle s^m s^n \rangle - \langle s^m \rangle \langle s^n \rangle$  with  $\langle s^m \rangle = w^{-1} \sum (s_i^m / \epsilon_i^2)$  and  $w = \sum \epsilon_i^{-2}$ .

The contribution due to multiple Coulomb scattering is approximately

$$\delta k_{\text{ms}} \approx \frac{(0.016)(\text{GeV}/c)z}{Lp\beta \cos^2 \lambda} \sqrt{\frac{L}{X_0}}, \quad (34.54)$$

where  $p$  = momentum (GeV/c)

$z$  = charge of incident particle in units of  $e$

$L$  = the total track length

$X_0$  = radiation length of the scattering medium (in units of length; the  $X_0$  defined elsewhere must be multiplied by density)

$\beta$  = the kinematic variable  $v/c$ .

More accurate approximations for multiple scattering may be found in the section on Passage of Particles Through Matter (Sec. 33 of this *Review*). The contribution to the curvature error is given approximately by  $\delta k_{\text{ms}} \approx 8 s_{\text{plane}}^{\text{rms}} / L^2$ , where  $s_{\text{plane}}^{\text{rms}}$  is defined there.

#### References:

1. T. Ferbel (ed.), *Experimental Techniques in High Energy Physics*, Addison-Wesley, Menlo Park, CA (1987).
2. K. Kleinknecht, *Detectors for Particle Radiation*, Cambridge University Press, Cambridge (1998).
3. G.F. Knoll, *Radiation Detection and Measurement*, 3rd edition, John Wiley & Sons, New York (1999).
4. D.R. Green, *The Physics of Particle Detectors*, Cambridge Monographs on Particle Physics, Nuclear Physics and Cosmology, Cambridge University Press, Cambridge (2000).
5. C. Leroy and P.-G. Rancoita, *Principles of Radiation Interaction in Matter and Detection*, World Scientific, Singapore (2004).
6. C. Grupen, *Particle Detectors*, Cambridge Monographs on Particle Physics, Nuclear Physics and Cosmology, Cambridge University Press (2008).
7. ICARUS Collab., ICARUS-TM/2001-09 [LGNS-EXP 13/89 add.2/01] (2001).
8. H. Spieler, IEEE Trans. **NS29**, 1142 (1982).
9. K. Arisaka, Nucl. Instrum. Methods **A442**, 80 (2000).
10. T. Hakamata (ed.), *Photomultiplier Tubes: Basics and Applications*, 3rd edition, Hamamatsu Photonics K.K., Hamamatsu (2006).
11. A. Braem *et al.*, Nucl. Instrum. Methods **A518**, 574 (2004).
12. R. Arnold *et al.*, Nucl. Instrum. Methods **A314**, 465 (1992).
13. P. Mangeot *et al.*, Nucl. Instrum. Methods **A216**, 79 (1983); R. Apsimon *et al.*, IEEE Trans. **NS33**, 112 (1986); R. Arnold *et al.*, Nucl. Instrum. Methods **A270**, 255 (1988); D. Aston *et al.*, Nucl. Instrum. Methods **A283**, 582 (1989).
14. J. Janesick, *Scientific charge-coupled devices*, SPIE Press, Bellingham, WA (2001).
15. R. Haitz *et al.*, J. Appl. Phys. **36**, 3123 (1965); R. McIntyre, IEEE Trans. Electron Devices **13**, 164 (1966); H. Dautet *et al.*, Applied Optics, **32**, 3894 (1993); Perkin-Elmer Optoelectronics, *Avalanche Photodiodes: A User's Guide*, (2003).
16. P. Buzhan *et al.*, Nucl. Instrum. Methods **A504**, 48 (2003); Z. Sadygov *et al.*, Nucl. Instrum. Methods **A504**, 301 (2003); V. Golovin and V. Saveliev, Nucl. Instrum. Methods **A518**, 560 (2004).
17. M. Landstrass *et al.*, Diam. & Rel. Matter, **2**, 1033 (1993); R. McKeag and R. Jackman, Diam. & Rel. Matter, **7**, 513 (1998); R. Brascia *et al.*, Phys. Stat. Sol. **199**, 113 (2003).
18. M. Petrov, M. Stapelbroek, and W. Kleinhaus, Appl. Phys. Lett. **51**, 406 (1987);

- M. Atac and M. Petrov, IEEE Trans. **NS36**, 163 (1989);  
M. Atac *et al.*, Nucl. Instrum. Methods **A314**, 56 (1994).
19. J.B. Birks, *The Theory and Practice of Scintillation Counting*, Pergamon, London (1964).
20. D. Clark, Nucl. Instrum. Methods **117**, 295 (1974).
21. J.B. Birks, Proc. Phys. Soc. **A64**, 874 (1951).
22. B. Bengtson and M. Moszynski, Nucl. Instrum. Methods **117**, 227 (1974);  
J. Bialkowski *et al.*, Nucl. Instrum. Methods **117**, 221 (1974).
23. C. P. Achenbach, arXiv:nucl-ex/0404008 (2004).
24. I.B. Berlman, *Handbook of Fluorescence Spectra of Aromatic Molecules*, 2nd edition, Academic Press, New York (1971).
25. F. Sauli (ed.), *Instrumentation in High Energy Physics*, World Scientific, Singapore (1992), see pp. 218–279 by C. Zorn.
26. T. Foerster, Ann. Phys. **2**, 55 (1948).
27. J.M. Fluornoy, Rad. Phys. and Chem. **41**, 389 (1993).
28. D. Horstman and U. Holm, Rad. Phys. and Chem. **41**, 395 (1993).
29. D. Blomker *et al.*, Nucl. Instrum. Methods **A311**, 505 (1992);  
J. Mainusch *et al.*, Nucl. Instrum. Methods **A312**, 451 (1992).
30. K.F. Johnson and R.L. Clough (eds.), *Proceedings of the International Conference on Radiation-Tolerant Plastic Scintillators and Detectors*, Rad. Phys. and Chem. **41** (1993).
31. S.R. Borenstein and R.C. Strand, IEEE Trans. **NS31**, 396 (1984).
32. P. Sonderegger, Nucl. Instrum. Methods **A257**, 523 (1987).
33. P. Achenbach *et al.*, Nucl. Instrum. Methods **A593**, 353 (2008).
34. C.M. Hawkes *et al.*, Nucl. Instrum. Methods **A292**, 329 (1990).
35. S.E. Derenzo, W.-S. Choong and W.W. Moses, Phys. Med. Biol. **59**, 3261 (2014).
36. C. Melcher and J. Schweitzer, Nucl. Instrum. Methods **A314**, 212 (1992).
37. D.W. Cooke *et al.*, J. Appl. Phys. **88**, 7360 (2000).
38. J.M. Chen *et al.*, IEEE Trans. **NS54**, 718 (2007);  
J.M. Chen *et al.*, IEEE Trans. **NS54**, 1319 (2007).
39. E.V.D. van Loef *et al.*, Nucl. Instrum. Methods **A486**, 254 (2002).
40. W. Drozdowski *et al.*, IEEE Trans. **NS55**, 1391 (2008).
41. M.S. Alekhin *et al.*, Appl. Phys. Lett. **102**, 161915 (2013).
42. C. Kuntner *et al.*, Nucl. Instrum. Methods **A493**, 131 (2002).
43. N. Akchurin *et al.*, Nucl. Instrum. Methods **A595**, 359 (2008).
44. A. Para, FERMILAB-CONF-11-519-CD (2011);  
H. Wenzel *et al.*, FERMILAB-PUB-11-531-CD-E (2011).
45. R.H. Mao, L.Y. Zhang and R.Y. Zhu, IEEE Trans. **NS59**, 2229 (2012).
46. W.W. Moses, W.-S. Choong and S.E. Derenzo, Acta Physica Polonica **B7**, 725 (2014).
47. R.H. Mao, L.Y. Zhang and R.Y. Zhu, IEEE Trans. **NS55**, 2425 (2008).
48. B.D. Rooney and J.D. Valentine, IEEE Trans. **NS44**, 509 (1997).
49. W.W. Moses *et al.*, IEEE Trans. **NS55**, 1049 (2008).
50. G. Gratta, H. Newman, and R.Y. Zhu, Ann. Rev. Nucl. and Part. Sci. **44**, 453 (1994).
51. R.Y. Zhu, Nucl. Instrum. Methods **A413**, 297 (1998).
52. R.Y. Zhu, Journal of Physics: Conference Series **587**, 012055 (2015).
53. G. Dissertori *et al.*, Nucl. Instrum. Methods **745**, 1 (2014), and references therein.
54. R. Chipaux *et al.*, *Proc. of the 8th Int. Conf. on Inorganic Scintillators (SCINT2005)*, 369 (2005).
55. The CMS Electromagnetic Calorimeter Group, JINST **5** P03010 (2010).
56. M.G. Albrow *et al.*, JINST **7**, P10027 (2012).
57. B. Aubert *et al.* [BaBar Collab.], Nucl. Instrum. Methods **A479**, 1 (2002).
58. A. Abashian *et al.*, Nucl. Instrum. Methods **A479**, 117 (2002).
59. M. Shiozawa *et al.* [Super-Kamiokande Collab.], Nucl. Instrum. Methods **A433**, 240 (1999).
60. *Proceedings of the International Workshops on Ring Imaging Cherenkov Detectors*, Nucl. Instrum. Methods **A343**, 1 (1993);  
Nucl. Instrum. Methods **A371**, 1 (1996);  
Nucl. Instrum. Methods **A433**, 1 (1999); Nucl. Instrum. Methods **A502**, 1 (2003);  
Nucl. Instrum. Methods **A553**, 1 (2005); Nucl. Instrum. Methods **A595**, 1 (2008);  
Nucl. Instrum. Methods **A639**, 1 (2011); Nucl. Instrum. Methods **A766**, 1 (2014).
61. J. Litt and R. Meunier, Ann. Rev. Nucl. Sci. **23**, 1 (1973).
62. D. Bartlett *et al.*, Nucl. Instrum. Methods **A260**, 55 (1987).
63. B. Ratcliff, Nucl. Instrum. Methods **A502**, 211 (2003).
64. W. Blum, W. Riegler, and L. Rolandi, *Particle Detection with Drift Chambers*, Springer-Verlag, Berlin (2008).
65. L.G. Christophorou, *Atomic and Molecular Radiation Physics*, John Wiley & Sons, Hoboken (1971);  
I.B. Smirnov, Nucl. Instrum. Methods **A554**, 474 (2005);  
J. Berkowitz, *Atomic and Molecular Photoabsorption*, Academic Press, Cambridge (2002);  
<http://pdg.lbl.gov/2007/AtomicNuclearProperties>.
66. H. Bichsel, Nucl. Instrum. Methods **A562**, 154 (2006).
67. H. Fischle *et al.*, Nucl. Instrum. Methods **A301**, 202 (1991).
68. <http://rjd.web.cern.ch/rjd/cgi-bin/cross>.
69. A. Peisert and F. Sauli, CERN 84-08 (1984).
70. S. Biagi, Nucl. Instrum. Methods **A421**, 234 (1999).
71. <http://consult.cern.ch/writeup/magboltz/>.
72. E. McDaniel and E. Mason, *The Mobility and Diffusion of Ions in Gases*, John Wiley & Sons, Hoboken (1973);  
G. Shultz *et al.*, Rev. Phys. Appl. **12**, 67 (1977).
73. G. Charpak *et al.*, Nucl. Instrum. Methods **A62**, 262 (1968).
74. G. Charpak and F. Sauli, Ann. Rev. Nucl. Sci. **34**, 285 (1984).
75. T. Ferbel (ed.), *Experimental Techniques in High Energy Physics*, Addison-Wesley, Menlo Park, CA (1987), see “Principles of Operation of Multiwire Proportional and Drift Chambers”.
76. G. Charpak *et al.*, Nucl. Instrum. Methods **A167**, 455 (1979).
77. A.H. Walenta *et al.*, Nucl. Instrum. Methods **A92**, 373 (1971).
78. A. Breskin *et al.*, Nucl. Instrum. Methods **A124**, 189 (1975).
79. A. Breskin *et al.*, Nucl. Instrum. Methods **A156**, 147 (1978).
80. R. Bouclier *et al.*, Nucl. Instrum. Methods **A265**, 78 (1988).
81. H. Drumm *et al.*, Nucl. Instrum. Methods **A176**, 333 (1980).
82. D.R. Nygren and J.N. Marx, Phys. Today **31N10**, 46 (1978).
83. <http://www.ansoft.com>.
84. P. Beringer *et al.*, Nucl. Instrum. Methods **A254**, 542 (1987).
85. J. Virdee, Phys. Reports **403**, 401 (2004).
86. H. Walenta, Phys. Scripta **23**, 354 (1981).
87. J. Va’vra, Nucl. Instrum. Methods **A515**, 1 (2003);  
M. Titov, arXiv:physics/0403055 (2004).
88. M. Aleksa *et al.*, Nucl. Instrum. Methods **A446**, 435 (2000).
89. F. Sauli and A. Sharma, Ann. Rev. Nucl. Part. Sci. **49**, 341 (1999).
90. A. Oed, Nucl. Instrum. Methods **A263**, 351 (1988);  
A. Barr *et al.*, Nucl. Phys. (Proc. Supp.) **B61**, 264 (1988).
91. Y. Bagaturia *et al.*, Nucl. Instrum. Methods **A490**, 223 (2002).
92. J. Benlloch *et al.*, IEEE Trans. **NS45**, 234 (1998).
93. Y. Giomataris, Nucl. Instrum. Methods **A419**, 239 (1998).
94. F. Sauli, Nucl. Instrum. Methods **A386**, 531 (1997);  
A. Bressan *et al.*, Nucl. Instrum. Methods **A425**, 262 (1999).
95. S. Bachmann *et al.*, Nucl. Instrum. Methods **A479**, 294 (2002);  
A. Bressan *et al.*, Nucl. Instrum. Methods **A424**, 321 (1999).
96. Y. Giomataris *et al.*, Nucl. Instrum. Methods **A376**, 29 (1996).
97. J. Derre *et al.*, Nucl. Instrum. Methods **A459**, 523 (2001);  
G. Charpak *et al.*, Nucl. Instrum. Methods **A478**, 26 (2002).
98. I. Giomataris *et al.*, Nucl. Instrum. Methods **A560**, 405 (2006).
99. S. Duarte Pinto *et al.*, IEEE NSS Conf. Record **N08-4**, 1426 (2008).
100. L. Periale *et al.*, Nucl. Instrum. Methods **A478**, 377 (2002);  
R. Chechik *et al.*, Nucl. Instrum. Methods **A535**, 303 (2004);  
A. Breskin *et al.*, Nucl. Instrum. Methods **A598**, 107 (2009).
101. A. Di Mauro *et al.*, Nucl. Instrum. Methods **A581**, 225 (2007).
102. R. Bellazzini *et al.*, Nucl. Instrum. Methods **A535**, 477 (2004);  
M. Campbell *et al.*, Nucl. Instrum. Methods **A540**, 295 (2005);  
A. Bamberger *et al.*, Nucl. Instrum. Methods **A573**, 361



- (2007);  
T.Kim *et al.*, Nucl. Instrum. Methods **A599**, 173 (2008).
103. M. Chefdeville *et al.*, Nucl. Instrum. Methods **A556**, 490 (2006).
  104. M. Titov, [arXiv:physics/0403055](https://arxiv.org/abs/physics/0403055) (2004).
  105. <http://rd51-public.web.cern.ch/RD51-Public>.
  106. J. Alme *et al.*, Nucl. Instrum. Methods **A622**, 316 (2010).
  107. N. Abgrall *et al.*, Nucl. Instrum. Methods **A637**, 25 (2011).
  108. ALICE Collab., ALICE-PUBLIC-2015-004 (2015).
  109. A.H. Walenta *et al.*, Nucl. Instrum. Methods **161**, 45 (1979).
  110. H. Aihara *et al.*, IEEE Trans. **NS30**, 63 (1983).
  111. X. Artru *et al.*, Phys. Rev. **D12**, 1289 (1975);  
G.M. Garibian *et al.*, Nucl. Instrum. Methods **125**, 133 (1975).
  112. B. Dolgoshein, Nucl. Instrum. Methods **A326**, 434 (1993).
  113. *TRDs for the Third Millenium: Proc. 2nd Workshop on Advanced Transition Radiation Detectors for Accelerator and Space Applications*, Nucl. Instrum. Methods **A522**, 1 (2004).
  114. *TRDs for the Third Millenium: Proc. 4th Workshop on Advanced Transition Radiation Detectors for Accelerator and Space Applications*, Nucl. Instrum. Methods **A706**, 1 (2013).
  115. B. Beischer *et al.*, Nucl. Instrum. Methods **A583**, 485 (2007).
  116. A. Adronic and J.P. Wessels, Nucl. Instrum. Methods **A666**, 130 (2012).
  117. M. Petris *et al.*, Nucl. Instrum. Methods **A714**, 17 (2007).
  118. J. Apostolakis *et al.*, Rad. Phys. and Chem. **78**, 859 (2009).
  119. R. Santonico and R. Cardarelli, Nucl. Instrum. Methods **A187**, 377 (1981).
  120. V.V. Parkhomchuk, Yu.N. Pestov, and N.V. Petrovykh, Nucl. Instrum. Methods **93**, 269 (1971).
  121. E. Cerron Zeballos *et al.*, Nucl. Instrum. Methods **A374**, 132 (1996).
  122. R. Cardarelli *et al.*, Nucl. Instrum. Methods **A333**, 399 (1993).
  123. P. Camarri *et al.*, Nucl. Instrum. Methods **A414**, 317 (1998).
  124. G. Aielli *et al.*, Nucl. Instrum. Methods **A508**, 6 (2003).
  125. R. Cardarelli *et al.*, Nucl. Instrum. Methods **A382**, 470 (1996).
  126. R. Cardarelli *et al.*, JINST **8**, P01003 (2013).
  127. G. Aielli *et al.*, JINST **9**, C09030 (2014).
  128. R. Santonico, JINST **9**, C11007 (2014).
  129. R. Santonico *et al.*, Nucl. Instrum. Methods **A661**, S2 (2012).
  130. M. Bedjidian *et al.*, JINST **6** P02001 (2011).
  131. P. Fonte *et al.*, Nucl. Instrum. Methods **A443**, 201 (2000).
  132. L. Paolozzi *et al.*, PoS(RPC2012)065 (2012).
  133. S. An *et al.*, Nucl. Instrum. Methods **A594**, 39 (2008).
  134. C. Iacobaeus *et al.*, Nucl. Instrum. Methods **A513**, 244 (2003).
  135. G. Aielli *et al.*, Nucl. Instrum. Methods **A456**, 82 (2000).
  136. G. Aielli *et al.*, IEEE Trans. **NS53**, 567 (2006).
  137. H. Sakai *et al.*, Nucl. Instrum. Methods **A484**, 153 (2002).
  138. R. Santonico, JINST **8**, P04023 (2013).
  139. L. Lopes *et al.*, Nucl. Instrum. Methods **A533**, 69 (2003).
  140. H. Spieler, *Semiconductor Detector Systems*, Oxford Univ. Press, Oxford (2005).
  141. F. Scholze *et al.*, Nucl. Instrum. Methods **A439**, 208 (2000).
  142. G. Lindström *et al.*, Nucl. Instrum. Methods **A465**, 60 (2001).
  143. C. Da Via *et al.*, Nucl. Instrum. Methods **A509**, 86 (2003).
  144. G. Kramberger *et al.*, Nucl. Instrum. Methods **A481**, 297 (2002).
  145. O. Krasel *et al.*, IEEE Trans. **NS51**, 3055 (2004).
  146. G. Lindström *et al.*, Nucl. Instrum. Methods **A426**, 1 (1999).
  147. A. Holmes-Siedle and L. Adams, *Handbook of Radiation Effects*, 2nd edition, Oxford Univ. Press, Oxford (2002).
  148. V. Radeka, IEEE Trans. **NS15**, 455 (1968);  
V. Radeka, IEEE Trans. **NS21**, 51 (1974).
  149. F.S. Goulding, Nucl. Instrum. Methods **100**, 493 (1972);  
F.S. Goulding and D.A. Landis, IEEE Trans. **NS29**, 1125 (1982).
  150. W.R. Nelson, H. Hirayama, and D.W.O. Rogers, SLAC-265 (1985).
  151. R. Brun *et al.*, CERN DD/EE/84-1 (1987).
  152. D. Hitlin *et al.*, Nucl. Instrum. Methods **137**, 225 (1976);  
See also W. J. Willis and V. Radeka, Nucl. Instrum. Methods **120**, 221 (1974), for a more detailed discussion.
  153. R. Wigmans, *Calorimetry: Energy Measurement in Particle Physics*, Inter. Series of Monographs on Phys. **107**, Clarendon Press, Oxford (2000).
  154. ATLAS Collab., CERN/LHCC 96-41 (1996).
  155. CMS Collab., CERN/LHCC 97-33 (1997).
  156. N. Akchurin *et al.*, Nucl. Instrum. Methods **A399**, 202 (1997).
  157. B. Aubert *et al.*, Nucl. Instrum. Methods **A321**, 467 (1992).
  158. A. Artamonov *et al.*, JINST **3**, P02010 (2008).
  159. F. Ariztizabal *et al.*, Nucl. Instrum. Methods **A349**, 384 (1994).
  160. S. Abdullin *et al.*, Eur. Phys. J. **C53**, 139 (2008).
  161. M. Romalli, J. Phys. Conf. Series **404** 012050 (2012);  
C. Adloff *et al.*, [arXiv:1306.3037](https://arxiv.org/abs/1306.3037) (2013).
  162. T.A. Gabriel *et al.*, Nucl. Instrum. Methods **A338**, 336 (1994).
  163. D.E. Groom, Nucl. Instrum. Methods **A572**, 633 (2007);  
Erratum: D.E. Groom, Nucl. Instrum. Methods **A593**, 638 (2008).
  164. N. Akchurin *et al.*, Nucl. Instrum. Methods **A408**, 380 (1998);  
An energy-independent analysis of these data is given in Ref. 163.
  165. P. Adragna *et al.*, Nucl. Instrum. Methods **A615**, 158 (2010).
  166. C.W. Fabjan *et al.*, Phys. Lett. **B60**, 105 (1975).
  167. C. Leroy, J. Sirois, and R. Wigmans, Nucl. Instrum. Methods **A252**, 4 (1986).
  168. J.E. Brau and T.A. Gabriel, Nucl. Instrum. Methods **A238**, 489 (1985);  
H. Brückmann and H. Kowalski, ZEUS Int. Note 86/026 DESY, Hamburg (1986);  
R. Wigmans, Nucl. Instrum. Methods **A259**, 389 (1987);  
R. Wigmans, Nucl. Instrum. Methods **A265**, 273 (1988).
  169. P. Mockett, SLAC-267, 335 (1987).
  170. R. Wigmans, *Proc. 7th Inter. Conf. on Calorimetry in High Energy Physics*, 182 World Scientific, River Edge, NJ, (1998);  
N. Akchurin *et al.*, Nucl. Instrum. Methods **A537**, 537 (2005).
  171. G. Drews *et al.*, Nucl. Instrum. Methods **A335**, 335 (1990).
  172. M. Holder *et al.*, Nucl. Instrum. Methods **151**, 69 (1978).
  173. R.K. Bock, T. Hansl-Kozanecka, and T.P. Shah, Nucl. Instrum. Methods **186**, 533 (1981);  
Y.A. Kulchitsky and V.B. Vinogradov, Nucl. Instrum. Methods **A455**, 499 (2000).
  174. D. Acosta *et al.*, Nucl. Instrum. Methods **A316**, 184 (1997).
  175. N. Akchurin, *et al.*, Nucl. Instrum. Methods **A735**, 120 (2013).
  176. W. Walkowiak, Nucl. Instrum. Methods **A449**, 288 (2000).
  177. L.S. Miller *et al.*, Phys. Rev. **166**, 871 (1968).
  178. E. Shibamura *et al.*, Nucl. Instrum. Methods **A316**, 184 (1975).
  179. K. Yoshino *et al.*, Phys. Rev. **A14**, 438 (1976).
  180. A.O. Allen *et al.*, NSRDS-NBS-58 (1976).
  181. P. Benetti *et al.*, Nucl. Instrum. Methods **A32**, 361 (1993).
  182. A.M. Kalinin *et al.*, ATLAS-LARG-NO-058 (1996).
  183. G. Danby *et al.*, Phys. Rev. Lett. **9**, 36 (1962).
  184. S. van der Meer, CERN 61-07 (1961).
  185. J.A. Formaggio and G.P. Zeller, Rev. Mod. Phys. **84**, 1307 (2013).
  186. H. Faissner *et al.*, Phys. Lett. **B68**, 377 (1977).
  187. F.J. Hasert *et al.*, Nucl. Phys. **B73**, 1 (1974).
  188. N. Armenise *et al.*, Phys. Lett. **B81**, 385 (1979).
  189. S.J. Barish *et al.*, Phys. Rev. **D16**, 3103 (1977).
  190. N.J. Baker *et al.*, Phys. Rev. **D23**, 2499 (1981).
  191. J.W. Chapman *et al.*, Phys. Rev. **D14**, 5 (1976).
  192. A.E. Asratian *et al.*, Phys. Lett. **79**, 497 (1978).
  193. Y. Fukuda *et al.*, Nucl. Instrum. Methods **A501**, 418 (2003).
  194. Y. Fukuda *et al.*, Phys. Rev. **D74**, 072003 (2006).
  195. A.A. Aguilar-Arevalo *et al.*, Nucl. Instrum. Methods **A599**, 28 (2009).
  196. A.A. Aguilar-Arevalo *et al.*, Phys. Rev. **D79**, 072002 (2009).
  197. C. Athanassopoulos *et al.*, Nucl. Instrum. Methods **A388**, 149 (1997).
  198. K. Hiraide *et al.*, Phys. Rev. **D78**, 112004 (2008).
  199. L. Aliaga *et al.*, Nucl. Instrum. Methods **A743**, 130 (2014).
  200. D. Ayres *et al.*, FERMILAB-DESIGN-2007-01 (2007).
  201. M. Holder *et al.*, Nucl. Instrum. Methods **148**, 203 (1978).

- 202. W.K. Sakumoto *et al.*, Nucl. Instrum. Methods **A294**, 179 (1990).
- 203. D.A. Harris *et al.*, Nucl. Instrum. Methods **A447**, 377 (2000).
- 204. I. Ambats *et al.*, FERMILAB-DESIGN-1998-02 (1998).
- 205. C. Rubbia, CERN-EP-INT-77-08 (1977).
- 206. S. Amerio *et al.*, Nucl. Instrum. Methods **A527**, 329 (2004).
- 207. C. Anderson *et al.*, JINST **7**, 10020 (2012).
- 208. H. Chen *et al.*, FERMILAB-PROPOSAL-0974 (2007).
- 209. D.H. Perkins, Nature **159**, 126 (1947).
- 210. N. Uhida *et al.*, Nucl. Instrum. Methods **224**, 50 (1984).
- 211. S. Aoki *et al.*, Nucl. Instrum. Methods **A447**, 361 (2000).
- 212. K. Kodama *et al.*, Nucl. Instrum. Methods **B93**, 340 (1994).
- 213. S. Aoki, Nucl. Instrum. Methods **A473**, 192 (2001).
- 214. T. Adam *et al.*, Nucl. Instrum. Methods **A577**, 523 (2007).
- 215. D. Di Ferdinando *et al.*, Radiat. Meas. **44**, 840 (2009).
- 216. R. Acquafredda *et al.*, New J. Phys. **8**, 303 (2006).
- 217. A.N. Diddens *et al.*, Nucl. Instrum. Methods **178**, 27 (1980).
- 218. D. Geiregat *et al.*, Nucl. Instrum. Methods **A325**, 92 (1993).
- 219. L.A. Ahrens *et al.*, Nucl. Instrum. Methods **A254**, 515 (1987).
- 220. G. Gidal, LBL-91 Suppl., Rev. (1985).
- 221. J. Altegoer *et al.*, Nucl. Instrum. Methods **A404**, 96 (1998).
- 222. K. Abe *et al.*, Nucl. Instrum. Methods **A659**, 106 (2011).
- 223. T. Taylor, Phys. Scripta **23**, 459 (1980).
- 224. A. Yamamoto, Nucl. Instrum. Methods **A494**, 255 (2003).
- 225. A. Yamamoto, Nucl. Instrum. Methods **A453**, 445 (2000).
- 226. R.L. Gluckstern, Nucl. Instrum. Methods **24**, 381 (1963).
- 227. V. Karimäki, Nucl. Instrum. Methods **A410**, 284 (1998).

## 35. PARTICLE DETECTORS FOR NON-ACCELERATOR PHYSICS

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### 35.1. Introduction

Non-accelerator experiments have become increasingly important in particle physics. These include classical cosmic ray experiments, neutrino oscillation measurements, and searches for double-beta decay, dark matter candidates, and magnetic monopoles. The experimental methods are sometimes those familiar at accelerators (plastic scintillators, drift chambers, TRD's, *etc.*) but there is also instrumentation either not found at accelerators or applied in a radically different way. Examples are atmospheric scintillation detectors (Fly's Eye), massive Cherenkov detectors (Super-Kamiokande, IceCube), ultracold solid state detectors (CDMS). And, except for the cosmic ray detectors, radiologically ultra-pure materials are required.

In this section, some more important detectors special to terrestrial non-accelerator experiments are discussed. Techniques used in both accelerator and non-accelerator experiments are described in Sec. 28, Particle Detectors at Accelerators, some of which have been modified to accommodate the non-accelerator nuances.

Space-based detectors also use some unique instrumentation, but these are beyond the present scope of RPP.

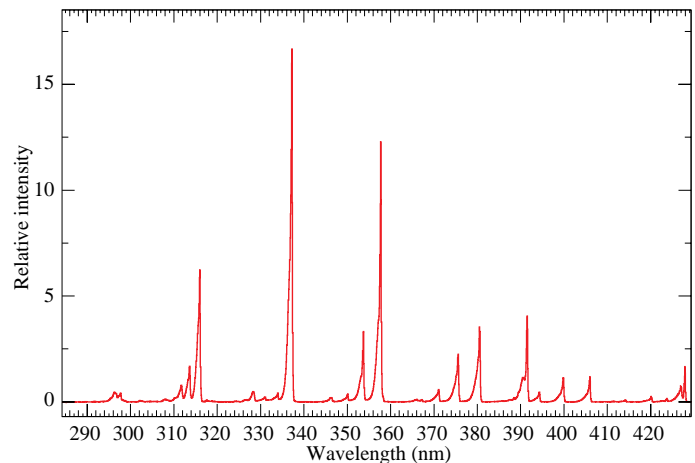
### 35.2. High-energy cosmic-ray hadron and gamma-ray detectors

#### 35.2.1. Atmospheric fluorescence detectors :

Revised August 2015 by L.R. Wiencke (Colorado School of Mines).

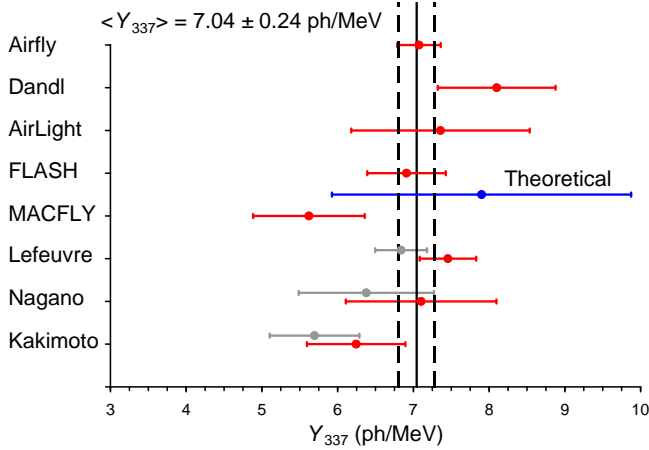
Cosmic-ray fluorescence detectors (FDs) use the atmosphere as a giant calorimeter to measure isotropic scintillation light that traces the development profiles of extensive air showers. An extensive air shower (EAS) is produced by the interactions of ultra high-energy ( $E > 10^{17}$  eV) subatomic particles in the stratosphere and upper troposphere. These are the highest energy particles known to exist. The amount of scintillation light generated is proportional to energy deposited in the atmosphere and nearly independent of the primary species. Experiments with FDs include the pioneering Fly's Eye [1], HiRes [2], the Telescope Array [3], and the Pierre Auger Observatory (Auger) [4]. The Auger FD also measures the time development of a class of atmospheric transient luminous events called "Elves" that are created in the ionosphere above some thunderstorms [5]. The proposed space based FD instrument [6] by the JEM-EUSO collaboration would look down on the earth's atmosphere from space to view a much larger area than ground based instruments.

The fluorescence light is emitted primarily between 290 and 430 nm (Fig. 35.1), when relativistic charged particles, primarily electrons and positrons, excite nitrogen molecules in air, resulting in transitions of the 1P and 2P systems. Reviews and references for the pioneering and recent laboratory measurements of fluorescence yield,  $Y(\lambda, P, T, u)$ , including dependence on wavelength ( $\lambda$ ), temperature ( $T$ ), pressure ( $p$ ), and humidity ( $u$ ) may be found in Refs. 7–9. The results of various experiments have been combined (Fig. 35.2) to obtain an absolute average and uncertainty for  $Y(337 \text{ nm}, 800 \text{ hPa}, 293 \text{ K}, \text{ dry air})$  of  $7.04 \pm 0.24 \text{ ph/MeV}$  after corrections for different electron beam energies and other factors. The units of ph/MeV correspond to the number of fluorescence photons produced per MeV of energy deposited in the atmosphere by the electromagnetic component of an EAS.



**Figure 35.1:** Measured fluorescence spectrum excited by 3 MeV electrons in dry air at 800 hPa and 293 K [11].

An FD element (telescope) consists of a non-tracking spherical mirror (3.5–13 m<sup>2</sup> and less than astronomical quality), a close-packed "camera" of photomultiplier tubes (PMTs) (for example, Hamamatsu R9508 or Photonis XP3062) near the focal plane, and a flash ADC readout system with a pulse and track-finding trigger scheme [10]. Simple reflector optics (12° × 16° degree field of view (FOV) on 256 PMTs) and Schmidt optics (30° × 30° FOV on 440 PMTs), including a correcting element, have been used. Segmented mirrors have been fabricated from slumped or slumped/polished glass with an anodized aluminium coating and from chemically anodized AlMgSiO<sub>5</sub> affixed to shaped aluminum. A broadband UV filter (custom fabricated or Schott MUG-6) reduces background light such as starlight, airglow, man-made light pollution, and airplane strobelights.



**Figure 35.2:** Fluorescence yield values and associated uncertainties at 337 nm ( $Y_{337}$ ) in dry air at 800 hPa and 293 K (Figure from [12]). The methodology and corrections that were applied to obtain the average and the uncertainty are discussed extensively in this reference. The vertical axis denotes different laboratory experiments that measured FY. The gray bars show three of the original measurements to illustrate the scale of the corrections applied.

At  $10^{20}$  eV, where the flux drops below 1 EAS/km<sup>2</sup>century, the aperture for an eye of adjacent FD telescopes that span the horizon can reach  $10^4$  km<sup>2</sup> sr. FD operation requires (nearly) moonless nights and clear atmospheric conditions, which imposes a duty cycle of about 10%. Arrangements of LEDs, calibrated diffuse sources [13], pulsed UV lasers [14], LIDARs\* and cloud monitors are used for photometric calibration, atmospheric calibration [15], and determination of exposure [16].

The EAS generates a track consistent with a light source moving at  $v = c$  across the FOV. The number of photons ( $N_\gamma$ ) as a function of atmospheric depth ( $X$ ) can be expressed as [8]

$$\frac{dN_\gamma}{dX} = \frac{dE_{\text{dep}}^{\text{tot}}}{dX} \int Y(\lambda, P, T, u) \cdot \tau_{\text{atm}}(\lambda, X) \cdot \varepsilon_{\text{FD}}(\lambda) d\lambda, \quad (35.1)$$

where  $\tau_{\text{atm}}(\lambda, X)$  is the atmospheric transmission, including wavelength ( $\lambda$ ) dependence, and  $\varepsilon_{\text{FD}}(\lambda)$  is the FD efficiency.  $\varepsilon_{\text{FD}}(\lambda)$  includes geometric factors and collection efficiency of the optics, quantum efficiency of the PMTs, and other throughput factors. The typical systematic uncertainties,  $\tau_{\text{atm}}$  (10%) and  $\varepsilon_{\text{FD}}$  (photometric calibration 10%), currently dominate the total reconstructed EAS energy uncertainty.  $\Delta E/E$  of 20% is possible, provided the geometric fit of the EAS axis is constrained typically by multi-eye stereo projection, or by timing from a colocated sparse array of surface detectors.

Analysis methods to reconstruct the EAS profile and deconvolute the contributions of re-scattered scintillation light, and direct and scattered Cherenkov light are described in [1] and more recently in [17]. The EAS energy is typically obtained by integrating over the Gaisser-Hillas function [18]

$$E_{\text{cal}} = \int_0^\infty w_{\text{max}} \left( \frac{X - X_0}{X_{\text{max}} - X_0} \right)^{(X_{\text{max}} - X_0)/\lambda} e^{(X_{\text{max}} - X)/\lambda} dX, \quad (35.2)$$

where  $X_{\text{max}}$  is the depth at which the shower reaches its maximum energy deposit  $w_{\text{max}}$ .  $X_0$  and  $\lambda$  are two shape parameters.

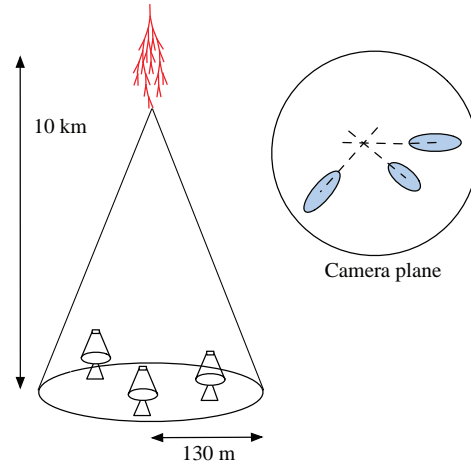
\* "LIDAR stands for "Light Detection and Ranging" and refers here to systems that measure atmospheric properties from the light scattered backwards from laser pulses directed into the sky.

### 35.2.2. Atmospheric Cherenkov telescopes for high-energy $\gamma$ -ray astronomy :

Revised November 2015 by J. Holder (Dept. of Physics and Astronomy & Bartol Research Inst., Univ. of Delaware).

A wide variety of astrophysical objects are now known to produce high-energy  $\gamma$ -ray photons. Leptonic or hadronic particles, accelerated to relativistic energies in the source, produce  $\gamma$ -rays typically through inverse Compton boosting of ambient photons or through the decay of neutral pions produced in hadronic interactions. At energies below  $\sim 30$  GeV,  $\gamma$ -ray emission can be efficiently detected using satellite or balloon-borne instrumentation, with an effective area approximately equal to the size of the detector (typically  $< 1$  m<sup>2</sup>). At higher energies, a technique with much larger effective collection area is required to measure astrophysical  $\gamma$ -ray fluxes, which decrease rapidly with increasing energy. Atmospheric Cherenkov detectors achieve effective collection areas of  $> 10^5$  m<sup>2</sup> by employing the Earth's atmosphere as an intrinsic part of the detection technique.

As described in Chapter 29, a hadronic cosmic ray or high energy  $\gamma$ -ray incident on the Earth's atmosphere triggers a particle cascade, or air shower. Relativistic charged particles in the cascade generate Cherenkov radiation, which is emitted along the shower direction, resulting in a light pool on the ground with a radius of  $\sim 130$  m. Cherenkov light is produced throughout the cascade development, with the maximum emission occurring when the number of particles in the cascade is largest, at an altitude of  $\sim 10$  km for primary energies of 100 GeV–1 TeV. Following absorption and scattering in the atmosphere, the Cherenkov light at ground level peaks at a wavelength,  $\lambda \approx 300$ –350 nm. The photon density is typically  $\sim 100$  photons/m<sup>2</sup> for a 1 TeV primary, arriving in a brief flash of a few nanoseconds duration. This Cherenkov pulse can be detected from any point within the light pool radius by using large reflecting surfaces to focus the Cherenkov light on to fast photon detectors (Fig. 35.3).



**Figure 35.3:** A schematic illustration of an imaging atmospheric Cherenkov telescope array. The primary particle initiates an air shower, resulting in a cone of Cherenkov radiation. Telescopes within the Cherenkov light pool record elliptical images; the intersection of the long axes of these images indicates the arrival direction of the primary, and hence the location of a  $\gamma$ -ray source in the sky.

Modern atmospheric Cherenkov telescopes, such as those built and operated by the VERITAS [19], H.E.S.S. [20] and MAGIC [21] collaborations, consist of large ( $> 100$  m<sup>2</sup>) segmented mirrors on steerable altitude-azimuth mounts. A camera made from an array of photosensors is placed at the focus of each mirror and used to record a Cherenkov image of each air shower. In these imaging atmospheric Cherenkov telescopes, single-anode photomultiplier tubes (PMTs) have traditionally been used (2048, in the case of H.E.S.S. II), but multi-anode PMTs and silicon devices now feature in more modern designs. The telescope cameras typically cover a field-of-view of  $3$ – $5^\circ$

in diameter. Images are recorded at kHz rates, the vast majority of which are due to showers with hadronic cosmic-ray primaries. The shape and orientation of the Cherenkov images are used to discriminate  $\gamma$ -ray photon events from this cosmic-ray background, and to reconstruct the photon energy and arrival direction.  $\gamma$ -ray images result from purely electromagnetic cascades and appear as narrow, elongated ellipses in the camera plane. The long axis of the ellipse corresponds to the vertical extension of the air shower, and it points back towards the source position in the field-of-view. If multiple telescopes are used to view the same shower (“stereoscopy”), the source position is simply the intersection point of the various image axes. Cosmic-ray primaries produce secondaries with large transverse momenta, which initiate sub-showers. Their images are consequently wider and less regular than those with  $\gamma$ -ray primaries and, since the original charged particle has been deflected by Galactic magnetic fields before reaching the Earth, the images have no preferred orientation.

The measurable differences in Cherenkov image orientation and morphology provide the background discrimination which makes ground-based  $\gamma$ -ray astronomy possible. For point-like sources, such as distant active galactic nuclei, modern instruments can reject over 99.999% of the triggered cosmic-ray events, while retaining up to 50% of the  $\gamma$ -ray population. In the case of spatially extended sources, such as Galactic supernova remnants, the background rejection is less efficient, but the technique can be used to produce  $\gamma$ -ray maps of the emission from the source. The angular resolution depends upon the number of telescopes which view the image and the energy of the primary  $\gamma$ -ray, but is typically less than  $0.1^\circ$  per event (68% containment radius) at energies above a few hundred GeV.

The total Cherenkov yield from the air shower is proportional to the energy of the primary particle. The image intensity, combined with the reconstructed distance of the shower core from each telescope, can therefore be used to estimate the primary energy. The energy resolution of this technique, also energy-dependent, is typically 15–20% at energies above a few hundred GeV. Energy spectra of  $\gamma$ -ray sources can be measured over a wide range, depending upon the instrument characteristics, source properties (flux, spectral slope, elevation angle, *etc.*), and exposure time: the H.E.S.S. measurement of the hard spectrum supernova remnant RX J1713.7–3946 extends to 100 TeV [22], for example, while pulsed emission from the Crab Pulsar has been detected at 25 GeV [23]. In general, peak sensitivity lies in the range from 100 GeV to a few TeV.

The first astrophysical source to be convincingly detected using the imaging atmospheric Cherenkov technique was the Crab Nebula [24], with an integral flux of  $2.1 \times 10^{-11}$  photons  $\text{cm}^{-2} \text{s}^{-1}$  above 1 TeV [25]. Modern imaging atmospheric Cherenkov telescopes have sensitivity sufficient to detect sources with less than 1% of the Crab Nebula flux in a few tens of hours. The TeV source catalog now consists of more than 160 sources (see e.g. Ref. 26). The majority of these were detected by scanning the Galactic plane from the southern hemisphere with the H.E.S.S. telescope array [27].

Major upgrades of the existing telescope arrays have recently been completed, including the addition of a 28 m diameter central telescope to H.E.S.S. (H.E.S.S. II). Development is also underway for the next generation instrument, the Cherenkov Telescope Array (CTA), which will consist of a northern and a southern hemisphere observatory, with a combined total of more than 100 telescopes [28]. Telescopes of three different sizes are planned, spread over an area of  $> 1 \text{ km}^2$ , providing wider energy coverage, improved angular and energy resolutions, and an order of magnitude improvement in sensitivity relative to existing imaging atmospheric Cherenkov telescopes. Baseline telescope designs are similar to existing devices, but exploit technological developments such as dual mirror optics and silicon photo-detectors.

### 35.3. Large neutrino detectors

#### 35.3.1. Deep liquid detectors for rare processes :

Revised August 2015 by K. Scholberg & C.W. Walter (Duke University)

Deep, large detectors for rare processes tend to be multi-purpose with physics reach that includes not only solar, reactor, supernova and atmospheric neutrinos, but also searches for baryon number violation, searches for exotic particles such as magnetic monopoles, and neutrino and cosmic-ray astrophysics in different energy regimes. The detectors may also serve as targets for long-baseline neutrino beams for neutrino oscillation physics studies. In general, detector design considerations can be divided into high- and low-energy regimes, for which background and event reconstruction issues differ. The high-energy regime, from about 100 MeV to a few hundred GeV, is relevant for proton decay searches, atmospheric neutrinos and high-energy astrophysical neutrinos. The low-energy regime (a few tens of MeV or less) is relevant for supernova, solar, reactor and geological neutrinos.

Large water Cherenkov and scintillator detectors (see Table 35.1) usually consist of a volume of transparent liquid viewed by photomultiplier tubes (PMTs) (see Sec. 34.2); the liquid serves as active target. PMT hit charges and times are recorded and digitized, and triggering is usually based on coincidence of PMT hits within a time window comparable to the detector’s light-crossing time. Because photosensors lining an inner surface represent a driving cost that scales as surface area, very large volumes can be used for comparatively reasonable cost. Some detectors are segmented into subvolumes individually viewed by PMTs, and may include other detector elements (*e.g.*, tracking detectors). Devices to increase light collection, *e.g.*, reflectors or waveshifter plates, may be employed. A common configuration is to have at least one concentric outer layer of liquid material separated from the inner part of the detector to serve as shielding against ambient background. If optically separated and instrumented with PMTs, an outer layer may also serve as an active veto against entering cosmic rays and other background events. The PMTs for large detectors typically range in size from 20 cm to 50 cm diameter, and typical quantum efficiencies are in the 20–25% range for scintillation and water-Cherenkov photons. PMTs with higher quantum efficiencies, 35% or higher, have recently become available. The active liquid volume requires purification and there may be continuous recirculation of liquid. For large homogeneous detectors, the event interaction vertex is determined using relative timing of PMT hits, and energy deposition is determined from the number of recorded photoelectrons. A “fiducial volume” is usually defined within the full detector volume, some distance away from the PMT array. Inside the fiducial volume, enough PMTs are illuminated per event that reconstruction is considered reliable, and furthermore, entering background from the enclosing walls is suppressed by a buffer of self-shielding. PMT and detector optical parameters are calibrated using laser, LED, or other light sources. Quality of event reconstruction typically depends on photoelectron yield, pixelization and timing.

Because in most cases one is searching for rare events, large detectors are usually sited underground to reduce cosmic-ray-related background (see Chapter 29). The minimum depth required varies according to the physics goals [29].

##### 35.3.1.1. Liquid scintillator detectors:

Past and current large underground detectors based on hydrocarbon scintillators include LVD, MACRO, Baksan, Borexino, KamLAND and SNO+. Experiments at nuclear reactors include CHOOZ, Double CHOOZ, Daya Bay, and RENO. Organic liquid scintillators (see Sec. 34.3.0) for large detectors are chosen for high light yield and attenuation length, good stability, compatibility with other detector materials, high flash point, low toxicity, appropriate density for mechanical stability, and low cost. They may be doped with waveshifters and stabilizing agents. Popular choices are pseudocumene (1,2,4-trimethylbenzene) with a few g/L of the PPO (2,5-diphenyloxazole) fluor, and linear alkylbenzene (LAB). In a typical detector configuration there will be active or passive regions of

undoped scintillator, non-scintillating mineral oil or water surrounding the inner neutrino target volume. A thin vessel or balloon made of nylon, acrylic or other material transparent to scintillation light may contain the inner target; if the scintillator is buoyant with respect to its buffer, ropes may hold the balloon in place. For phototube surface coverages in the 20–40% range, yields in the few hundreds of photoelectrons per MeV of energy deposition can be obtained. Typical energy resolution is about  $7\%/\sqrt{E(\text{MeV})}$ , and typical position reconstruction resolution is a few tens of cm at  $\sim 1$  MeV, scaling as  $\sim N^{-1/2}$ , where  $N$  is the number of photoelectrons detected.

Fiducialization and tagging can reduce background. One can also dissolve neutrinoless double beta decay ( $0\nu\beta\beta$ ) isotopes in scintillator. This has been realized by KamLAND-Zen, which deployed a 1.5 m-radius balloon containing enriched Xe dissolved in scintillator inside KamLAND, and  $^{130}\text{Te}$  is planned for SNO+. Although for this approach, energy resolution is poor compared to other  $0\nu\beta\beta$  search experiments, the quantity of isotope can be so large that the kinematic signature of  $0\nu\beta\beta$  would be visible as a clear feature in the spectrum.

**Table 35.1:** Properties of large detectors for rare processes. If total target mass is divided into large submodules, the number of subdetectors is indicated in parentheses.

Detector	Mass, kton (modules)	PMTs (diameter, cm)	$\xi$	p.e./MeV	Dates
Baksan	0.33, scint (3150)	1/module (15)	segmented	40	1980–
MACRO	0.56, scint (476)	2-4/module (20)	segmented	18	1989–2000
LVD	1, scint. (840)	3/module (15)	segmented	15	1992–
KamLAND	0.41 <sup>f</sup> , scint	1325(43)+554(51)*	34%	460	2002–
Borexino	0.1 <sup>f</sup> , scint	2212 (20)	30%	500	2007–
SNO+	0.78, scint	9438 (20)	54%	400–900	2016 (exp.)
CHOOZ	0.005, scint (Gd)	192 (20)	15%	130	1997–1998
Double Chooz	0.017, scint (Gd)(2)	534/module (20)	13%	180	2011–
Daya Bay	0.160, scint (Gd)(8)	192/module (20)	5.6% <sup>†</sup>	100	2011–
RENO	0.032, scint (Gd)(2)	342/module (25)	12.6%	100	2011–
IMB-1	3.3 <sup>f</sup> , H <sub>2</sub> O	2048 (12.5)	1%	0.25	1982–1985
IMB-2	3.3 <sup>f</sup> , H <sub>2</sub> O	2048 (20)	4.5%	1.1	1987–1990
Kam I	0.88/0.78 <sup>f</sup> , H <sub>2</sub> O	1000/948 (50)	20%	3.4	1983–1985
Kam II	1.04 <sup>f</sup> , H <sub>2</sub> O	948 (50)	20%	3.4	1986–1990
Kam III	1.04 <sup>f</sup> , H <sub>2</sub> O	948 (50)	20% <sup>‡</sup>	4.3	1990–1995
SK I	22.5 <sup>f</sup> , H <sub>2</sub> O	11146 (50)	39%	6	1996–2001
SK II	22.5 <sup>f</sup> , H <sub>2</sub> O	5182 (50)	19%	3	2002–2005
SK III+	22.5 <sup>f</sup> , H <sub>2</sub> O	11129 (50)	39%	6	2006–
SNO	1, D <sub>2</sub> O/1.7, H <sub>2</sub> O	9438 (20)	31% <sup>§</sup>	9	1999–2006

<sup>f</sup> indicates typical fiducial mass used for data analysis; this may vary by physics topic.

\* Measurements made before 2003 only considered data from the 43 cm PMTs.

<sup>†</sup> The effective Daya Bay coverage is 12% with top and bottom reflectors.

<sup>‡</sup> The effective Kamiokande III coverage was 25% with light collectors.

<sup>§</sup> The effective SNO coverage was 54% with light collectors.

Shallow detectors for reactor neutrino oscillation experiments require excellent muon veto capabilities. For  $\bar{\nu}_e$  detection via inverse beta decay on free protons,  $\bar{\nu}_e + p \rightarrow n + e^+$ , the neutron is captured by a proton on a  $\sim 180$   $\mu\text{s}$  timescale, resulting in a 2.2 MeV  $\gamma$  ray, observable by Compton scattering and which can be used as a tag in coincidence with the positron signal. The positron annihilation  $\gamma$  rays may also contribute. Inverse beta decay tagging may be improved by addition of Gd at  $\sim 0.1\%$  by mass, which for natural isotope abundance has a  $\sim 49,000$  barn cross-section for neutron capture (in contrast to the 0.3 barn cross-section for capture on free protons). Gd capture takes  $\sim 30$   $\mu\text{s}$ , and is followed by a cascade of  $\gamma$  rays adding up to about 8 MeV. Gadolinium doping of scintillator requires specialized formulation to ensure adequate attenuation length and stability.

Scintillation detectors have an advantage over water Cherenkov detectors in the lack of Cherenkov threshold and the high light yield. However, scintillation light emission is nearly isotropic, and therefore directional capabilities are relatively weak. Liquid scintillator is especially suitable for detection of low-energy events. Radioactive backgrounds are a serious issue, and include long-lived cosmogenics. To go below a few MeV, very careful selection of materials and purification of the scintillator is required (see Sec. 35.6).

### 35.3.1.2. Water Cherenkov detectors:

Very large imaging water detectors reconstruct ten-meter-scale Cherenkov rings produced by charged particles (see Sec. 34.5.0). The first such large detectors were IMB and Kamiokande. The only currently existing instance of this class of detector, with fiducial volume of 22.5 kton and total mass of 50 kton, is Super-Kamiokande (Super-K). For volumes of this scale, absorption and scattering of Cherenkov light are non-negligible, and a wavelength-dependent factor  $\exp(-d/L(\lambda))$  (where  $d$  is the distance from emission to the sensor and  $L(\lambda)$  is the attenuation length of the medium) must be included in the integral of Eq. (34.5) for the photoelectron yield. Attenuation lengths on the order of 100 meters have been achieved.

Cherenkov detectors are excellent electromagnetic calorimeters, and the number of Cherenkov photons produced by an  $e/\gamma$  is nearly proportional to its kinetic energy. For massive particles, the number of photons produced is also related to the energy, but not linearly. For any type of particle, the *visible energy*  $E_{\text{vis}}$  is defined as the energy of an electron which would produce the same number of Cherenkov photons. The number of collected photoelectrons depends on the scattering and attenuation in the water along with the photocathode coverage, quantum efficiency and the optical parameters

of any external light collection systems or protective material surrounding them. Event-by-event corrections are made for geometry and attenuation. For a typical case, in water  $N_{p.e.} \sim 15 \xi E_{vis}(\text{MeV})$ , where  $\xi$  is the effective fractional photosensor coverage. Cherenkov photoelectron yield per MeV of energy is relatively small compared to that for scintillator, *e.g.*,  $\sim 6$  pe/MeV for Super-K with a PMT surface coverage of  $\sim 40\%$ . In spite of light yield and Cherenkov threshold issues, the intrinsic directionality of Cherenkov light allows individual particle tracks to be reconstructed. Vertex and direction fits are performed using PMT hit charges and times, requiring that the hit pattern be consistent with a Cherenkov ring.

High-energy ( $\sim 100$  MeV or more) neutrinos from the atmosphere or beams interact with nucleons; for the nucleons bound inside the  $^{16}\text{O}$  nucleus, nuclear effects must be considered both at the interaction and as the particles leave the nucleus. Various event topologies can be distinguished by their timing and fit patterns, and by presence or absence of light in a veto. “Fully-contained” events are those for which the neutrino interaction final state particles do not leave the inner part of the detector; these have their energies relatively well measured. Neutrino interactions for which the lepton is not contained in the inner detector sample have higher-energy parent neutrino energy distributions. For example, in “partially-contained” events, the neutrino interacts inside the inner part of the detector but the lepton (almost always a muon, since only muons are penetrating) exits. “Upward-going muons” can arise from neutrinos which interact in the rock below the detector and create muons which enter the detector and either stop, or go all the way through (entering downward-going muons cannot be distinguished from cosmic rays). At high energies, multi-photoelectron hits are likely and the charge collected by each PMT (rather than the number of PMTs firing) must be used; this degrades the energy resolution to approximately  $2\%/\sqrt{\xi E_{vis}(\text{GeV})}$ . The absolute energy scale in this regime can be known to  $\sim 2\text{--}3\%$  using cosmic-ray muon energy deposition, Michel electrons and  $\pi^0$  from atmospheric neutrino interactions. Typical vertex resolutions for GeV energies are a few tens of cm [30]. Angular resolution for determination of the direction of a charged particle track is a few degrees. For a neutrino interaction, because some final-state particles are usually below Cherenkov threshold, knowledge of direction of the incoming neutrino direction itself is generally worse than that of the lepton direction, and dependent on neutrino energy.

Multiple particles in an interaction (so long as they are above Cherenkov threshold) may be reconstructed, allowing for the exclusive reconstruction of final states. In searches for proton decay, multiple particles can be kinematically reconstructed to form a decaying nucleon. High-quality particle identification is also possible:  $\gamma$  rays and electrons shower, and electrons scatter, which results in fuzzy rings, whereas muons, pions and protons make sharp rings. These patterns can be quantitatively separated with high reliability using maximum likelihood methods [31]. A  $e/\mu$  misidentification probability of  $\sim 0.4\%/\xi$  in the sub-GeV range is consistent with the performance of several experiments for  $4\% < \xi < 40\%$ . Sources of background for high energy interactions include misidentified cosmic muons and anomalous light patterns when the PMTs sometimes “flash” and emit photons themselves. The latter class of events can be removed using its distinctive PMT signal patterns, which may be repeated. More information about high energy event selection and reconstruction may be found in reference [32].

In spite of the fairly low light yield, large water Cherenkov detectors may be employed for reconstructing low-energy events, down to *e.g.*  $\sim 4\text{--}5$  MeV for Super-K [33]. Low-energy neutrino interactions of solar neutrinos in water are predominantly elastic scattering off atomic electrons; single electron events are then reconstructed. At solar neutrino energies, the visible energy resolution ( $\sim 30\%/\sqrt{\xi E_{vis}(\text{MeV})}$ ) is about 20% worse than photoelectron counting statistics would imply. Using an electron LINAC and/or nuclear sources, approximately 0.5% determination of the absolute energy scale has been achieved at solar neutrino energies. Angular resolution is limited by multiple scattering in this energy regime ( $25\text{--}30^\circ$ ). At these energies, radioactive backgrounds become a dominant issue. These backgrounds include radon in the water itself or emanated from detector materials, and  $\gamma$  rays from the rock and

detector materials. In the few to few tens of MeV range, radioactive products of cosmic-ray-muon-induced spallation are troublesome, and are removed by proximity in time and space to preceding muons, at some cost in dead time. Gadolinium doping using 0.2%  $\text{Gd}_2(\text{SO}_4)_3$  is planned for Super-K to improve selection of low-energy  $\bar{\nu}_e$  and other events with accompanying neutrons [34].

The Sudbury Neutrino Observatory (SNO) detector [35] is the only instance of a large heavy water detector and deserves mention here. In addition to an outer 1.7 kton of light water, SNO contained 1 kton of  $\text{D}_2\text{O}$ , giving it unique sensitivity to neutrino neutral current ( $\nu_x + d \rightarrow \nu_x + p + n$ ), and charged current ( $\nu_e + d \rightarrow p + p + e^-$ ) deuteron breakup reactions. The neutrons were detected in three ways: In the first phase, via the reaction  $n + d \rightarrow t + \gamma + 6.25$  MeV; Cherenkov radiation from electrons Compton-scattered by the  $\gamma$  rays was observed. In the second phase, NaCl was dissolved in the water.  $^{35}\text{Cl}$  captures neutrons,  $n + ^{35}\text{Cl} \rightarrow ^{36}\text{Cl} + \gamma + 8.6$  MeV. The  $\gamma$  rays were observed via Compton scattering. In a final phase, specialized low-background  $^3\text{He}$  counters (“neutral current detectors” or NCDs) were deployed in the detector. These counters detected neutrons via  $n + ^3\text{He} \rightarrow p + t + 0.76$  MeV; ionization charge from energy loss of the products was recorded in proportional counters.

### 35.3.2. Neutrino telescopes :

Revised Nov. 2015 by Ch. Spiering (DESY/Zeuthen) and U.F. Katz (Univ. Erlangen)

The primary goal of neutrino telescopes (NTs) is the detection of astrophysical neutrinos, in particularly those which are expected to accompany the production of high-energy cosmic rays in astrophysical accelerators. NTs in addition address a variety of other fundamental physics issues like indirect search for dark matter, study of neutrino oscillations, search for exotic particles like magnetic monopoles or study of cosmic rays and their interactions [36,37,38].

NTs are large-volume arrays of “optical modules” (OMs) installed in open transparent media like water or ice, at depths that completely block the daylight. The OMs record the Cherenkov light induced by charged secondary particles produced in reactions of high-energy neutrinos in or around the instrumented volume. The neutrino energy,  $E_\nu$ , and direction can be reconstructed from the hit pattern recorded. NTs typically target an energy range  $E_\nu \gtrsim 100$  GeV; sensitivity to lower energies is achieved in dedicated setups with denser instrumentation.

In detecting cosmic neutrinos, three sources of backgrounds have to be considered: (i) *atmospheric neutrinos* from cosmic-ray interactions in the atmosphere, which can be separated from cosmic neutrinos only on a statistical basis; (ii) down-going punch-through *atmospheric muons* from cosmic-ray interactions, which are suppressed by several orders of magnitude with respect to the ground level due to the large detector depths. They can be further reduced by selecting upward-going or high-energy muons or by self-veto methods sensitive to the muon entering the detector; (iii) random backgrounds due to photomultiplier (PMT) dark counts,  $^{40}\text{K}$  decays (mainly in sea water) or bioluminescence (only water), which impact adversely on event recognition and reconstruction. Note that atmospheric neutrinos and muons allow for investigating neutrino oscillations and cosmic ray anisotropies, respectively.

Recently, it has become obvious that a precise measurement of the energy-zenith-distribution of atmospheric neutrinos may allow for determining the neutrino mass hierarchy by exploiting matter-induced oscillation effects in the Earth.

Neutrinos can interact with target nucleons  $N$  through charged current ( $\bar{\nu}_\ell N \rightarrow \ell^\mp X$ , CC) or neutral current ( $\bar{\nu}_\ell N \rightarrow \bar{\nu}_\ell X$ , NC) processes. A CC reaction of a  $\bar{\nu}_\mu$  produces a muon track and a hadronic particle cascade, whereas all NC reactions and CC reactions of  $\bar{\nu}_e$  produce particle cascades only. CC interactions of  $\bar{\nu}_\tau$  can have either signature, depending on the  $\tau$  decay mode. In most astrophysical models, neutrinos are produced through the  $\pi/K \rightarrow \mu \rightarrow e$  decay chain, *i.e.*, with a flavour ratio  $\nu_e : \nu_\mu : \nu_\tau \approx 1 : 2 : 0$ . For sources outside the solar system, neutrino oscillations turn this ratio to  $\nu_e : \nu_\mu : \nu_\tau \approx 1 : 1 : 1$  upon arrival on Earth.

The total neutrino-nucleon cross section is about  $10^{-35} \text{ cm}^2$  at

$E_\nu = 1$  TeV and rises roughly linearly with  $E_\nu$  below this energy and as  $E_\nu^{0.3-0.5}$  above, flattening out towards high energies. The CC:NC cross-section ratio is about 2:1. At energies above some TeV, neutrino absorption in the Earth becomes noticeable; for vertically upward-moving neutrinos (zenith angle  $\theta = 180^\circ$ ), the survival probability is 74 (27, < 2)% for 10 (100, 1000) TeV. On average, between 50% (65%) and 75% of  $E_\nu$  is transferred to the final-state lepton in neutrino (antineutrino) reactions between 100 GeV and 10 PeV.

NC reactions or  $\tau$  decays may carry away significant “invisible” energy. Above 100 TeV, the directional reconstruction accuracy of cascades is 10–15 degrees in polar ice and about 2 degrees in water, the difference being due to the inhomogeneity of the ice and the stronger light scattering in ice. These features, together with the small background of atmospheric  $\bar{\nu}_e$  and  $\bar{\nu}_\tau$  events, makes the cascade channel particularly interesting for searches for a diffuse, high-energy excess of extraterrestrial over atmospheric neutrinos. In water,

**Table 35.2:** Past, present and planned neutrino telescope projects and their main parameters. The milestone years give the times of project start, of first data taking with partial configurations, of detector completion, and of project termination. Projects with first data expected past 2020 are indicated in italics. The size refers to the largest instrumented volume reached during the project development. See [38] for references to the different projects where unspecified.

Experiment, Milestones	Medium, Location	Size [km <sup>3</sup> ]	Remarks
DUMAND, 1978/–/–/1995	Pacific/Hawaii		Terminated due to technical/funding problems
NT-200 1980/1993/1998/–	Lake Baikal	10 <sup>−4</sup>	First proof of principle
GVD [39] 2012/2015/–/–	Lake Baikal	0.5–1.5	High-energy $\nu$ astronomy, first cluster installed
NESTOR 1991/–/–/–	Med. Sea		2004 data taking with prototype
NEMO 1998/–/–/–	Med. Sea		R&D project, prototype tests
AMANDA 1990/1996/2000/2009	Ice/South Pole	0.015	First deep-ice neutrino telescope
ANTARES 1997/2006/2008/2016	Med. Sea	0.010	First deep-sea neutrino telescope
IceCube 2001/2005/2010/–	Ice/South Pole	1.0	First km <sup>3</sup> -sized detector
PINGU [40] 2014/–/–/–	Ice/South Pole	0.003	Planned low-energy extension of IceCube
<i>IceCube-Gen2</i> [41] 2014/–/–/–	Ice/South Pole	5–10	Planned high-energy extension
KM3NeT/ARCA 2013/(2017)/–/–	Med. Sea	1–2	First construction phase started
KM3NeT/ORCA 2014/(2017)/–/–	Med. Sea	0.003	Low-energy configuration for neutrino mass hierarchy
<i>KM3NeT Phase 3</i> 2013/–/–/–	Med. Sea	3–6	6 building blocks + ORCA

The final-state lepton follows the initial neutrino direction with a RMS mismatch angle  $\langle\phi_{\nu\ell}\rangle \approx 1.5^\circ/\sqrt{E_\nu [\text{TeV}]}$ , indicating the intrinsic kinematic limit to the angular resolution of NTs. For CC  $\bar{\nu}_\mu$  reactions at energies above about 10 TeV, the angular resolution is dominated by the muon reconstruction accuracy of a few times  $0.1^\circ$  at most. For muon energies  $E_\mu \gtrsim 1$  TeV, the increasing light emission due to radiative processes allows for reconstructing  $E_\mu$  from the measured  $dE_\mu/dx$  with an accuracy of  $\sigma(\log E_\mu) \approx 0.3$ ; at lower energies,  $E_\mu$  can be estimated from the length of the muon track if it is contained in the detector. These properties make CC  $\bar{\nu}_\mu$  reactions the prime channel for the identification of individual astrophysical neutrino sources.

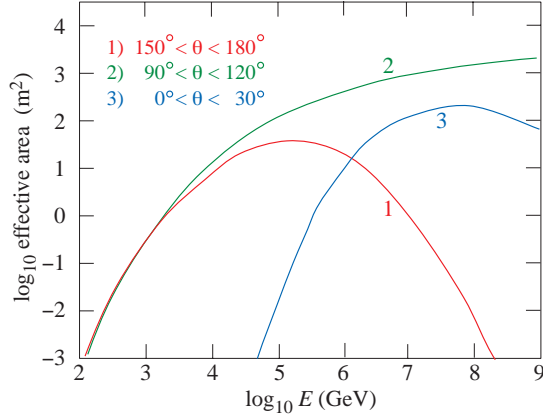
Hadronic and electromagnetic particle cascades at the relevant energies are 5–20 m long, *i.e.*, short compared to typical distances between OMs. The total amount of Cherenkov light provides a direct measurement of the cascade energy with an accuracy of about 20% at energies above 10 TeV and 10% beyond 100 TeV for events contained in the instrumented volume. Neutrino flavour and reaction mechanism can, however, hardly be determined and neutrinos from

cascade events can also be used for the search for point sources of cosmic neutrinos. The inferior angular accuracy compared to muon tracks, however, leads to a higher number of background events per source from atmospheric neutrinos.

The detection efficiency of a NT is quantified by its effective area, *e.g.*, the fictitious area for which the full incoming neutrino flux would be recorded (see Fig. 35.4). The increase with  $E_\nu$  is due to the rise of neutrino cross section and muon range, while neutrino absorption in the Earth causes the decrease at large  $\theta$ . Identification of downward-going neutrinos requires strong cuts against atmospheric muons, hence the cut-off towards low  $E_\nu$ . Due to the small cross section, the effective area is many orders of magnitude smaller than the geometrical dimension of the detector; a  $\bar{\nu}_\mu$  with 1 TeV can, *e.g.*, be detected with a probability of the order  $10^{-6}$  if the telescope is on its path.

Detection of upward going muons makes the effective volume of the detector much larger than its geometrical volume. The method, however, is only sensitive to CC  $\bar{\nu}_\mu$  interactions and cannot be





**Figure 35.4:** Effective  $\bar{\nu}_\mu$  area for IceCube as an example of a cubic-kilometre NT, as a function of neutrino energy for three intervals of the zenith angle  $\theta$ . The effective areas shown here correspond to a specific event selection for point source searches.

extended much more than 5–10 degrees above the geometric horizon, where the background of atmospheric muons becomes prohibitive. Alternatively, one can select events that start inside the instrumented volume. In contrast to neutrinos, incoming muons generate early hits in the outer layers of the detector. Such a veto-based event selection is sensitive to neutrinos of all flavours from all directions, albeit with a reduced effective volume since a part of the instrumented volume is sacrificed for the veto. The muon veto also rejects down-going atmospheric neutrinos that typically are accompanied by muons in the same air shower and thus reduces the atmospheric-neutrino background. Actually, the breakthrough in detecting high-energy cosmic neutrinos has been achieved with this technique.

Note that the fields of view of NTs at the South Pole and in the Northern hemisphere are complementary for each reaction channel and neutrino energy.

### 35.3.2.1. Properties of media:

The efficiency and quality of event reconstruction depend strongly on the optical properties (absorption and scattering length, intrinsic optical activity) of the medium in the spectral range of alkali photocathodes (300–550 nm). Large absorption lengths result in a better light collection, large scattering lengths in superior angular resolution. Deep-sea sites typically have effective scattering lengths of  $> 100$  m and, at their peak transparency around 450 nm, absorption lengths of 50–65 m. The absorption length for Lake Baikal is 22–24 m. The properties of South Polar ice vary strongly with depth; at the peak transparency wave length (400 nm), the scattering length is between 5 and 75 m and the absorption length between 15 and 250 m, with the best values in the depth region 2200–2450 m and the worst ones in the layer 1950–2100 m.

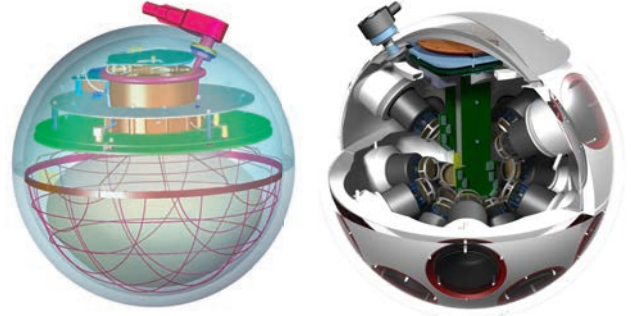
Noise rates measured by 25 cm PMTs in deep polar ice are about 0.5 kHz per PMT and almost entirely due to radioactivity in the OM components. The corresponding rates in sea water are typically 60 kHz, mostly due to  $^{40}\text{K}$  decays. Bioluminescence activity can locally cause rates on the MHz scale for seconds; the frequency and intensity of such “bursts” depends strongly on the sea current, the season, the geographic location, and the detector geometry. Experience from ANTARES shows that these backgrounds are manageable without a major loss of efficiency or experimental resolution.

### 35.3.2.2. Technical realisation:

**Optical modules (OMs) and PMTs:** An OM is a pressure-tight glass sphere housing one or several PMTs with a time resolution in the nanosecond range, and in most cases also electronics for control, HV generation, operation of calibration LEDs, time synchronisation and signal digitisation.

Hybrid PMTs with 37 cm diameter have been used for NT-200, conventional hemispheric PMTs for AMANDA (20 cm) and for ANTARES, IceCube and Baikal-GVD (25 cm). A novel concept has been chosen for KM3NeT. The OMs (43 cm) are equipped with 31

PMTs (7.5 cm), plus control, calibration and digitisation electronics. The main advantages are that (i) the overall photocathode area exceeds that of a 25 cm PMT by more than a factor of 3; (ii) the individual readout of the PMTs results in a very good separation between one- and two-photoelectron signals which is essential for online data filtering and random background suppression; (iii) the hit pattern on an OM provides directional information; (iv) no mu-metal shielding against the Earth magnetic field is required. Figure 35.5 shows the OM designs of IceCube and KM3NeT.



**Figure 35.5:** Schematic views of the digital OMs of IceCube (left) and KM3NeT (right).

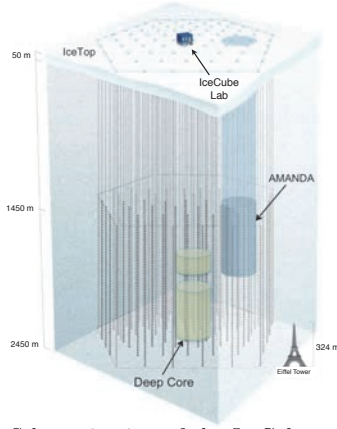
**Readout and data filtering:** In current NTs the PMT data are digitised in situ, for ANTARES and Baikal-GVD in special electronics containers close to the OMs, for IceCube and KM3NeT inside the OMs. For IceCube, data are transmitted via electrical cables of up to 3.3 km length, depending on the location of the strings and the depth of the OMs; for ANTARES, KM3NeT and Baikal-GVD optical fibre connections have been chosen (several 10 km for the first two and 4 km for GVD).

The full digitised waveforms of the IceCube OMs are transmitted to the surface for pulses appearing in local coincidences on a string; for other pulses, only time and charge information is provided. For ANTARES (time and charge) and KM3NeT (time over threshold), all PMT signals above an adjustable noise threshold are sent to shore.

The raw data are subsequently processed on online computer farms, where multiplicity and topology-driven filter algorithms are applied to select event candidates. The filter output data rate is about 10 GByte/day for ANTARES and of the order 1 TByte/day for IceCube (100 GByte/day transferred via satellite) and KM3NeT.

**Calibration:** For efficient event recognition and reconstruction, the OM timing must be synchronised at the few-nanosecond level and the OM positions and orientations must be known to a few 10 cm and a few degrees, respectively. Time calibration is achieved by sending synchronisation signals to the OM electronics and also by light calibration signals emitted by LED or laser flashers emitted in situ at known times (ANTARES, KM3NeT). Precise position calibration is achieved by measuring the travel time of light calibration signals sent from OM to OM (IceCube) or acoustic signals sent from transducers at the sea floor to receivers on the detector strings (ANTARES, KM3NeT, Baikal-GVD). Absolute pointing and angular resolution can be determined by measuring the “shadow of the moon” (*i.e.*, the directional depletion of muons generated in cosmic-ray interactions). IceCube has shown that both are below  $1^\circ$ , confirming MC calculations which indicate a precision of  $\approx 0.5^\circ$  for energies above 10 TeV. For KM3NeT, simulations indicate that sub-degree precision in the absolute pointing can be reached within a few weeks of operation.

**Detector configurations:** IceCube (see Fig. 35.6) consists of 5160 Digital OMs (DOMs) installed on 86 strings at depths of 1450 to 2450 m in the Antarctic ice; except for the DeepCore region, string distances are 125 m and vertical distances between OMs 17 m. 324 further DOMs are installed in IceTop, an array of detector stations on the ice surface above the strings. DeepCore is a high-density sub-array at large depths (*i.e.*, in the best ice layer) at the centre of IceCube.



**Figure 35.6:** Schematic view of the IceCube neutrino observatory comprising the deep-ice detector including its nested dense part DeepCore, and the surface air shower array IceTop. The IceCube Lab houses data acquisition electronics and the computer farm for online processing. Operation of AMANDA was terminated in 2009.

The NT200 detector in Lake Baikal at a depth of 1100 m consists of 8 strings attached to an umbrella-like frame, with 12 pairs of OMs per string. The diameter of the instrumented volume is 42 m, its height 70 m. The Baikal collaboration has installed the first cluster of a future cubic-kilometre array. A first phase, covering a volume of about  $0.4 \text{ km}^3$ , will consist of 12 clusters, each with 192–288 OMs at 8 strings; its completion is scheduled for 2020. A next stage could comprise 27 clusters and cover up to  $1.5 \text{ km}^3$ .

ANTARES comprises 12 strings with lateral separations of 60–70 m, each carrying 25 triplets of OMs at vertical distances of 14.5 m. The OMs are located at depths 2.1–2.4 km, starting 100 m above the sea floor. A further string carries devices for calibration and environmental monitoring. A system to investigate the feasibility of acoustic neutrino detection is also implemented.

KM3NeT will consist of building blocks of 115 strings each, with 18 OMs per string. Prototype operations have successfully verified the KM3NeT technology [42]. Phase 2.0 of KM3NeT aims to demonstrate two separate detector arrangements, ARCA and ORCA. ARCA (Astroparticle Research with Cosmics in the Abyss) will search for high-energy astrophysical neutrinos using a sparse arrangement of OMs, with vertical separations of 36 m and a lateral separation between strings of 90 m. ORCA (Oscillation Research with Cosmics in the Abyss) intends to measure the neutrino mass hierarchy using a densely-packed arrangement, with 6–12 m vertical and 20 m lateral separations.

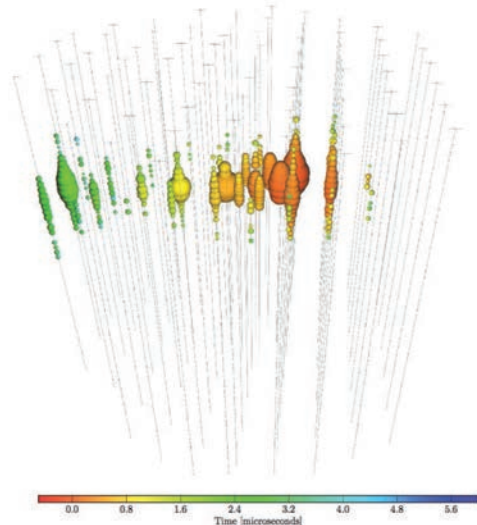
A first installation phase of ARCA near Capo Passero, East of Sicily and of ORCA near Toulon has started in 2014 and comprises 24 (7) ARCA (ORCA) strings to be deployed by the end of 2016. Completion of the full three blocks is expected for 2020.

### 35.3.2.3. Results:

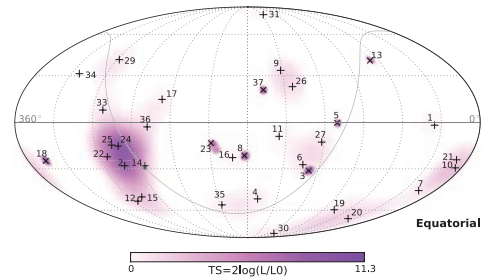
Atmospheric neutrino fluxes have been precisely measured with AMANDA and ANTARES ( $\bar{\nu}_\mu$ ) and with IceCube ( $\bar{\nu}_\mu, \bar{\nu}_e$ ); the results are in agreement with predicted spectra. No astrophysical point sources have been identified yet, and no indications of neutrino fluxes from dark matter annihilations or of exotic phenomena have been found (see [38] and references therein). IceCube has furthermore reported an energy-dependent anisotropy of cosmic-ray induced muons.

In 2013, an excess of track and cascade events between 30 TeV and 1 PeV above background expectations was reported by IceCube; this analysis used the data taken in 2010 and 2011 and for the first time employed containment conditions and an atmospheric muon veto for suppression of down-going atmospheric neutrinos (High-Energy Starting Event analysis, HESE). A display of one of the selected events is shown in Fig. 35.7. The observed excess reached a significance of  $5.7\sigma$  in a subsequent analysis of 3 years

of data [43] and cannot be explained by atmospheric neutrinos and misidentified atmospheric muons alone. Some clustering of the HESE events close to the Galactic Centre was observed (see Fig. 35.8). The hypothesis that this low-significance excess could be due to a point source with a spectral index of  $\geq 2$  was constrained by an analysis of ANTARES data looking at lower energies and with superior pointing to the same sky region [44]. Meanwhile the energy range of the IceCube HESE analysis has been extended down to 1 TeV and the high-energy excess confirmed; also, events with through-going muons showed a corresponding excess of cosmic origin. In [45], the various analyses have been combined. Assuming the cosmic neutrino flux to be isotropic, flavour-symmetric and  $\nu$ - $\bar{\nu}$ -symmetric at Earth, the all-flavour spectrum is well described by a power law with normalization  $6.7^{+1.1}_{-1.2} \times 10^{-18} \text{ GeV}^{-1} \text{ s}^{-1} \text{ sr}^{-1} \text{ cm}^{-2}$  at 100 TeV and a spectral index  $-2.50 \pm 0.09$  for energies between 25 TeV and 2.8 PeV. A spectral index of  $-2$ , an often quoted benchmark value, is disfavoured with a significance of  $3.8\sigma$ .



**Figure 35.7:** Event display of one of the starting-track events (event no. 5 in Fig. 35.8) from [43]. The deposited energy is 70 TeV, the colour code indicates the signal timing (red: early; green: late).



**Figure 35.8:** Arrival directions of 37 candidate events for cosmic neutrinos in equatorial coordinates (from [43]). Shower-like events (median angular resolution 15 degrees) are marked with + and those containing muon tracks degree) with x. Approximately 40% of the events are expected to originate from atmospheric backgrounds. The grey curve denotes the galactic plane and the grey dot the galactic centre. Colours show the test statistic for a point source clustering test at each location, with no significant clustering observed.

At lower energies, down to 10 GeV, IceCube/DeepCore and ANTARES have identified clear signals of oscillations of atmospheric neutrinos. The closely spaced OMs of DeepCore allow selecting a very pure sample of low-energy  $\bar{\nu}_\mu$  (6–56 GeV) that produce upward

moving muons inside the detector. The neutrino energy is determined from the energy of the hadronic shower at the vertex and the muon range. Fits to the energy/zenith-dependent deficit of muon neutrinos provide constraints on the oscillation parameters  $\sin^2 \theta_{23}$  and  $\Delta m_{23}^2$  (see the update of fig. 14.6 in the 2014 PDG).

See [46] and [47] for summaries of recent results of IceCube and ANTARES, respectively.

#### 35.3.2.4. Plans beyond 2020:

It is planned to extend the sensitivity of IceCube towards both lower and higher energies. A substantially denser instrumentation of a sub-volume of DeepCore would lead to an  $E_\nu$  threshold for neutrino detection of a few GeV. This project (*Phased IceCube Next Generation Upgrade*, PINGU) [40] primarily aims at measuring the neutrino mass hierarchy. For higher energies, a large-volume extension called IceCube-Gen2, combined with a powerful surface veto, is discussed [41]. More information on the future extensions of GVD and KM3NeT are given above and in Table 35.2.

#### 35.3.3. Coherent radio Cherenkov radiation detectors :

Revised August 2015 by S.R. Klein (LBNL/UC Berkeley)

Radio-frequency detectors are an attractive way to search for coherent Cherenkov radiation from showers produced from interactions of ultra-high energy cosmic neutrinos. These neutrinos are produced when protons with energy  $E > 4 \times 10^{19}$  eV interact with cosmic microwave background radiation (CMB) photons and are excited to a  $\Delta^+$  resonance. The subsequent  $\Delta^+ \rightarrow n\pi^+$  decay leads to the production of neutrinos with energies above  $10^{18}$  eV [48]. Neutrinos are the only long-range probe of the ultra-high energy cosmos, because protons, heavier nuclei and photons with energies above  $5 \times 10^{19}$  eV are limited to ranges of less than 100 Mpc by interactions with CMB photons and early starlight.

To detect this cosmic neutrino signal (of at least a few events per year, assuming that ultra-high energy cosmic-rays are protons) requires a detector of about  $100 \text{ km}^3$  in volume, made out of a non-conducting solid (or potentially liquid) medium, with a long absorption length for radio waves. The huge target volumes require that this be a commonly available natural material. A dense medium would be ideal to reduce the detector volume, but, unfortunately, the available natural media are of only moderate density. Optical Cherenkov and acoustical detectors are limited by a short ( $\sim 300$  m) attenuation length [49] so would require a prohibitive number of sensors. Radio-detection is the only current approach that can scale to this volume. The two commonly used media are Antarctic (or Greenland) ice and the lunar regolith [50]. Table 35.3 compares the characteristics of these different media, including several possible ice locations.

**Table 35.3:** Characteristics of different detection media for radio-Cherenkov signals. The attenuation length is at a frequency of 300 MHz; the Greenland figure is extrapolated upward from the 75 MHz measurements. The Moon and ice have similar Cherenkov angles because they have similar indices of refraction.

Medium	Density	Cherenkov Ang.	Cutoff Freq.	Atten. Length
Lunar Regolith	2.5 g/cm <sup>3</sup>	56°	3.0 GHz	9m/f(GHz) [50]
Antarctic Ice (South Pole)	0.92 g/cm <sup>3</sup>	56°	1.15 GHz	900 m [54]
Ross Ice Shelf	0.92 g/cm <sup>3</sup>	56°	1.15 GHz	406 m [55]
Greenland	0.92 g/cm <sup>3</sup>	56°	1.15 GHz	1022 m [56]

Electromagnetic and hadronic showers produce radio pulses via the Askaryan effect [51], as discussed in Sec. 33. The shower contains more electrons than positrons, leading to coherent emission.

High-frequency radiation is concentrated around the Cherenkov angle. On the cone, the electric field strength at a frequency  $f$  from an electromagnetic shower from a  $\nu_e$  may be roughly parameterized

as [52]

$$E_{\text{Ch}}(\text{V/MHz}) = 2.53 \times 10^{-7} \frac{E_\nu}{1 \text{ TeV}} \frac{f}{f_c} \left[ \frac{1}{1 + (f/f_c)^{1.44}} \right]. \quad (35.3)$$

The electric field strength increases linearly with frequency, up to a cut-off  $f_c$ , which is set by the transverse size of the shower [53]; the maximum wavelength is roughly the Moliere radius divided by  $\cos(\theta_C)$  where  $\theta_C$  is the Cherenkov angle. Some examples are given in Table 35.3.

Near  $f_c$ , radiation is narrowly concentrated around the Cherenkov angle [53]. At lower frequencies, the limited length of the emitting region leads to a diffractive broadening in emission angle with respect to the Cherenkov cone. The electric field from Eq. (35.3) is reduced by [52],

$$\frac{E}{E_{\text{Ch}}} = \exp \left( -\frac{1}{2} \frac{(\theta - \theta_C)^2}{(2.20 \times [1 \text{ GHz}/f])^2} \right), \quad (35.4)$$

At very low frequencies, the distribution is nearly isotropic.

Along the Cherenkov cone, the initial pulse width is  $\approx 1$  nsec, but it may be broadened by dispersion as it propagates, particularly for signals traversing the ionosphere. As long as the dispersion can be compensated for, a large bandwidth detector is the most sensitive. Spectral information can be used to reject background, and to help reconstruct the neutrino direction, because the cutoff frequency depends on the observation angle (with respect to the Cherenkov cone).

The detection threshold is determined by the distance to the antenna and the noise characteristics of the detector. Since the signal is a radio wave, its amplitude decreases as  $1/R$ , plus absorption in the intervening medium. Once anthropogenic noise is eliminated (not always easy), the main noise source is thermal noise. This can be reduced with careful design; locating a detector in cold ice also helps. Other potential backgrounds include cosmic-ray air showers, charge generated by blowing snow, and lightning.

The field is linearly proportional to the neutrino energy, so the power (field strength squared) is proportional to the square of the neutrino energy. For an antenna located in the detection medium, the typical threshold is around  $10^{17}$  eV; for stand-off (remote sensing) detectors, the threshold rises roughly linearly with the distance. These thresholds can be reduced significantly by using directional antennas and/or combining the signals from multiple antennas using beam-forming techniques. Experiments have used both approaches to reduce trigger-level noise, or to reject background at the analysis level. Optimally, the threshold will drop linearly with the square root of number of antennas, since the signal adds linearly while the background is added with random phases.

The signal is linearly polarized in the plane perpendicular to the neutrino direction. This polarization is an important check that any observed signal is indeed coherent Cherenkov radiation. Polarization measurements can be used to help reconstruct the neutrino direction.

At energies above  $10^{20}$  eV, the Landau-Pomeranchuk-Migdal effect significant spreads out electromagnetic showers, producing what are

effectively subshowers with significant separation. In this regime, the radio emission becomes even more concentrated around the Cherenkov cone, and then, at higher energies the emission begins to vary event-by-event. Because of this, many of the experiments that study higher energy neutrinos focus on the hadronic shower from the struck nucleus. This contains on average only about 20% of the energy, but with large fluctuations. It is of interest for very high energy searches (far above  $10^{20}$  eV) because it is much less subject to the LPM effects.

Radio detectors have observed cosmic-ray air showers in the atmosphere. The physics of radio-wave generation in air showers is more complex because there is a large contribution due to charge separation as electrons and positrons are bent in different directions as they propagate, leading to a growing charge dipole (transverse current) [57]. This time-varying transverse current emits radiation, spread over the transverse size of the shower. Since the radiating particles are moving relativistically downward, a ground-based observer sees a Lorentz contracted pulse which can have frequency components reaching the GHz range, limited by the thickness of the particle shower. There is also a contribution from geosynchrotron radiation, as  $e^\pm$  are bent in the same field [57]. The Askaryan effect is relatively small compared to these other sources. Experiments optimized for  $\nu$  detection can also detect air showers [58], which presents a potential background. Magnetic monopoles would also emit radio waves, and neutrino experiments have also set monopole flux limits [59].

#### 35.3.4. The Moon as a target :

Because of its large size and non-conducting regolith, and the availability of large radio-telescopes, the Moon is an attractive target [60]; some of the lunar experiments are listed in Table 35.4. Conventional radio-telescopes are reasonably well suited to lunar neutrino searches, with natural beam widths not too dissimilar from the size of the Moon. Still, there are experimental challenges. The composition of the lunar regolith is not well known, and the attenuation length for radio waves must be estimated. The big limitation of lunar experiments is that the 240,000 km target-antenna separation leads to neutrino energy thresholds above  $10^{20}$  eV.

**Table 35.4:** Experiments that have set limits on neutrino interactions in the Moon; current limits are shown in Fig. 1 of [50], with Lunaska (2015) from [68].

Experiment	Year	Dish Size	Frequency	Bandwidth	Obs. Time
Parkes	1995	64 m	1425 MHz	500 MHz	10 hrs
Glue	1999+	70 m, 34 m	2200 MHz	40-150 MHz	120 hrs
NuMoon	2008	11×25 m	115–180 MHz	—	50 hrs
Lunaska	2008	3× 22 m	1200–1800 MHz	—	6 nights
Lunaska	2015	64	1200-1500 MHz	300 MHz	127 hours
Resun	2008	4× 25 m	1450 MHz	50 MHz	45 hours

The effective volume probed by experiments depends on the geometry, which itself depends on the frequency range used. At low frequencies, radiation is relatively isotropic, so signals can be detected from most of the Moon's surface, for most angles of incidence. Also, radio signals penetrate more deeply at low frequencies, so the volume is larger than at shorter wavelengths. At higher frequencies, the electric field strength is higher, but radiation is concentrated near the Cherenkov angle. So, high-frequency experiments are only sensitive for a narrow range of geometries where the neutrino interacts near the Moon's limb with the Cherenkov cone pointed toward the Earth. Because of the stronger electric fields at high frequencies, these experiments are sensitive to lower energy neutrinos, albeit with a smaller effective volume, which gives them a lower flux sensitivity.

With modern technology, it is increasingly viable to search over very broad frequency ranges [61]. One technical challenge is due to dispersion (frequency dependent time delays) in the atmosphere. Dispersion can be largely removed with a de-dispersion filter, using either analog circuitry or post-collection digital processing.

Anthropogenic backgrounds are a major concern for ultra-high energy neutrino experiments. Lunar experiments use different techniques to reduce this background. Some experiments use multiple antennas, separated by at least hundreds of meters; by requiring a coincidence within a small time window, anthropogenic noise can be rejected. If the timing is good enough, beam-forming techniques can be used to further reduce the background. An alternative approach is to use beam forming with multiple feed antennas viewing a single reflector, to ensure that the signal points back to the moon.

These efforts have considerable scope for expansion. In the near future, several large radio detector arrays should reach significantly lower limits. The LOFAR array is beginning to take data with 36 detector clusters spread over Northwest Europe. In the longer term, the Square Kilometer Array (SKA) with  $1 \text{ km}^2$  effective area will push thresholds down to near  $10^{20}$  eV.

#### 35.3.5. Ice-based detectors :

Lower energy thresholds require a smaller antenna-target separation. Natural ice is an attractive medium for this, with attenuation lengths over 300 m. The attenuation length varies with the frequency and ice temperature, with higher attenuation in warmer ice. Table 35.3 lists some measurements of radio attenuation.

Although the ice is mostly uniform, the top  $\approx 100$  m of Antarctic ice, the 'firn,' contains a gradual transition from packed snow at the surface (typical surface density  $0.35 \text{ g/cm}^3$ ) to solid ice (density  $0.92 \text{ g/cm}^3$ ) below [62]. The index of refraction depends linearly on the density, so radio waves curve downward in the firn. This bending reduces the effectiveness of surface or aerial antennas. The thickness of the firn varies with location; it is thicker in central Antarctica than in the coastal ice sheets. For above-ice observations, it is also necessary to consider the surface roughness of the ice, which can affect signals as they transition from the ice to the atmosphere.

There are two types of Antarctic neutrino experiments. In one class, antennas mounted on scientific balloons observe the ice from above. The ANITA experiment is one example. It made two flights around Antarctica, floating at an altitude around 35 km [63]. Its 40 (32 in the first flight) dual-polarization horn antennas scanned the surrounding ice, out to the horizon (650 km away). Because of the small angle of incidence, ANITA could make use of polarization information;  $\nu$  signals should be vertically polarized, while most background from cosmic-ray air showers is expected to be horizontally polarized.

Because of the significant source-detector separation, ANITA is most sensitive at energies above  $10^{19}$  eV, above the peak of the GZK neutrino spectrum. As with the lunar experiments, ANITA had to contend with anthropogenic backgrounds. The ANITA collaboration uses their multiple antennas as a phased array to achieve good pointing accuracy, and used that to remove all apparent signals that pointed toward known or suspected areas of human habitation. By using the several-meter separation between antennas, they achieved a pointing accuracy of  $0.2\text{--}0.4^\circ$  in elevation, and  $0.5\text{--}1.1^\circ$  in azimuth. ANITA has set the most stringent limits on GZK neutrinos to date.

The proposed EVA experiment will use a portion of a fixed-shape balloon as a large parabolic radio antenna. Because of the large antenna surface, they hope to achieve threshold around  $10^{17}$  eV.

Other ice based experiments use antennas located within the active volume, allowing them to reach thresholds around  $10^{17}$  eV. This approach was pioneered by the RICE experiment, which buried 18 half-wave dipole antennas in holes drilled for AMANDA [64] at the South Pole, at depths from 100 to 300 m. The hardware was sensitive from 200 MHz to 1 GHz. Each antenna fed an in-situ preamplifier which transmitted the signals to surface digitizing electronics.

Three groups are prototyping detectors, with the goal of a detector with a  $\sim 100 \text{ km}^3$  active volume. For all three concepts, the hardware is modular, so the detector volume scales roughly linearly with the available funding. The Askaryan Radio Array (ARA) is located at the South Pole [65], while the Antarctic Ross Ice Shelf Antenna Neutrino Array (ARIANNA) is on the Ross Ice Shelf [66]. The Greenland Neutrino Observatory (GNO) collaboration is proposing a detector near the U.S. Summit Station in Greenland [67].

All of the in-ice experiments use multiple antennas, with varying degrees of connection. ARIANNA and ARA use the timing between multiple antennas in a single station to determine the arrival direction. At larger distance scales, such as between ARA and ARIANNA stations, the relative timing uncertainty is larger, and the antennas are treated as effectively independent, with independent triggers, and the data is only combined in offline analyses.

One big difference between the experiments is the depth of their antennas. ARA buries their antennas up to 200 m deep in the ice, to avoid the firn. Because of the refraction, a surface antenna cannot ‘see’ a signal from a near-surface interaction some distance away. However, drilling holes has costs, and the limited hole diameter (15 cm in ARA) requires compromises between antenna design (particularly for horizontally polarized waves), mechanical support, power and communications. In contrast, ARIANNA places their antennas in shallow, near-surface holes. This greatly simplifies deployment and avoids limitations on antenna design, but at a cost of reduced sensitivity to neutrino interactions near the surface. Because ARIANNA is at a green-field site, anthropogenic noise is much less of a problem.

The current ARA proposal, ARA-37 [65], calls for an array of 37 stations, each consisting of 16 embedded antennas. ARA will detect signals from 150 to 850 MHz for vertical polarization, and 250 MHz to 850 MHz for horizontal polarization. ARA plans to use bicone antennas for vertical polarization, and quad-slotted cylinders for horizontal polarization. The collaboration uses notch filters and surface veto antennas to eliminate most anthropogenic noise, and vetos events when aircraft are in the area, or weather balloons are being launched.

ARIANNA is in Moore’s Bay, on the Ross Ice Shelf, where  $\approx 575$  m of ice sits atop the Ross Sea [69]. The site was chosen because the ice-seawater interface is smooth there, so the interface acts as a mirror for radio waves. The major advantage of this approach is that ARIANNA is sensitive to downward going neutrinos, and should be able to see more of the Cherenkov cone for horizontal neutrinos. One disadvantage of the site is that the ice is warmer, so the radio attenuation length will be shorter. Each ARIANNA station will use six or eight log-periodic dipole antennas, pointing downward; two upward-pointing antennas will be used to veto cosmic-ray air showers and other backgrounds [66]. The multiple antennas allow for single-station directional and polarization measurements. The ARIANNA site is about 110 km from McMurdo station, and is shielded by Minna Bluff.

All three experiments share some significant challenges. Solar cells provide power during the 6-month summer, but the winter is a challenge. To date, wind power has not worked well, due to a combination of the low temperatures, harsh environment, and limited wind speed. ARA and GNO can run through the winter, at a cost of stringing long cables between stations and the base, but ARIANNA will likely only take data for 7 months/year. Also, because of its latitude, GNO could run for a large fraction of the year using solar power.

### 35.4. Large time-projection chambers for rare event detection

Written Nov. 2015 by T. Shutt (SLAC).

Rare event searches require detectors that combine large target masses and low levels of radioactivity, and that are located deep underground to eliminate cosmic-ray related backgrounds. Past and present efforts include searches for the scattering of particle dark matter, neutrinoless double beta decay, and the measurement of solar neutrinos, while next generation experiments will also probe coherent scattering of solar, atmospheric and diffuse supernova background neutrinos. Large time project chambers (TPCs), adapted from particle collider experiments, have emerged as a leading technology for these efforts. Events are measured in a central region confined by a field cage and usually filled with a liquid noble element target. Ionized electrons are drifted (in the  $z$  direction) to an anode region by use of electrode grids and field shaping rings, where their magnitude and

$x - y$  location is measured. In low background TPCs, scintillation generated at the initial event site is also measured, and the time difference between this prompt signal and the later-arriving charge signal gives the event location in  $z$  for a known electron drift speed. Thus, 3D imaging is achieved in a monolithic central volume. Noble elements have relatively high light yields (comparable to or exceeding the best inorganic scintillators), and the charge signal can be amplified by multiplication or electroluminescence. Radioactive backgrounds are distinguished by event imaging, the separate measurements of charge and light, and scintillation pulse shape. For recent reviews of noble element detectors, see [70,71].

Methods for achieving very low radioactive backgrounds are discussed in general in section 34.6. The basic architecture of large TPCs is very favorable for this application because gas or liquid targets can be relatively easily purified, while the generally more radioactive readout and support materials are confined to the periphery. The 3D imaging of the TPC then allows self shielding in the target material, which is quite powerful when the target is large compared to mean scattering lengths of  $\sim$  MeV neutrons and gammas from radioactivity ( $\sim 10$  cm in LXe, for example). The use of higher density targets (i.e., liquid instead of gas and/or higher mass elements) maximizes the ratio of target to surrounding material mass. The TPC geometry allows highly hermetic external shielding, with recent experiments using large water shields, in some cases enhanced with an active liquid scintillator layer.

In noble element targets, all non-noble impurities are readily removed (e.g., by chemical reaction in a commercial getter) so that only radioactive noble isotopes are a significant background concern. Xe, Ne and He have no long lived radioactive isotopes (apart from the  $^{136}\text{Xe}$ , discussed below). Kr has  $\sim 1$  MBq/kg of the beta emitter  $^{85}\text{Kr}$  created by nuclear fuel reprocessing, making it unusable as a target, while the 1 Bq/kg level of the beta emitter  $^{39}\text{Ar}$  is a nuisance for Ar-based experiments. Both of these can be backgrounds in other target materials, as can Rn emanating from detector components. Relatively low background materials are available for most of the structures surrounding the central target, with the exception of radioactive glasses and ceramics usually present in PMTs, feedthroughs and electrical components. Very low background PMTs with synthetic quartz windows, developed over the last decade, have been a key enabling technology for dark matter searches. These are not yet low enough in background for some double beta decay searches, which use radio-clean Si-based photon detectors.

An important technical challenge in liquid detectors is achieving the high voltages needed for electron drift and measurement. Quench gases which stabilize charge gain and speed electron transport in wire chambers cannot be used, since these absorb scintillation light (and also suppress charge extraction in dual-phase detectors, discussed below). At low energies (e.g., in a dark matter search) it is also important to suppress low-level emission of electrons and associated photons. Drift of electrons over meter scales with minimal loss from attachment on trace levels of dissolved impurities (e.g.,  $\text{O}_2$ ) has so far required continuous circulating purification. The relatively slow readout due to  $\sim$  msec/m drift speeds is not a major pile-up concern in low background experiments.

#### 35.4.1. Dark matter and other low energy signals :

A major goal of low background experiments is detection of WIMP (Weakly Interacting Massive Particle) dark matter through scattering on nuclei in a terrestrial detector (for a recent review, see [72]). Energy transfers are generally small, a few tens of keV at most. Liquid noble TPCs distinguish nuclear recoils (NR) from dark matter from the usually dominant background of electron recoils (ER) from gamma rays and beta decays by requiring single scatters and based on their charge to light ratio or scintillation pulse shape, as described below. Neutrons are a NR background, but can be recognized in a large imaging TPC if they multiply scatter. To detect small charge signals, a dual phase technique is used wherein electrons from interactions in the liquid target are drifted to the liquid surface and extracted with high field ( $\sim 5$  kV/cm) into the gas phase leading to an amplified electroluminescence signal measured by an array of PMTs located just above. (Both charge multiplication and electroluminescence are



possible in liquid, but have seen little use because the signals have very broad dispersion). This technique readily measures single electrons with  $\sim \text{cm } x-y$  resolution. The sides of the chamber are lined with highly (diffusively) reflective PTFE, and a second PMT array is located below the active volume to maximize the sensitivity to the initial scintillation signal.

The microscopic processes leading to signals in liquid nobles are complex. Energy deposited by an event generates pairs of free electron and ions, and also atoms in their lowest excited state. These rapidly form excimers which de-excite by emitting light. Excimers arise in both triplet and singlet states which have the same energy but different decay times. In an event track, some fraction of electrons recombine with ions, while the rest escape and are measured. Recombination leads to further excimer formation and hence more scintillation photons. Finally, some part of the energy is lost as heat - a small fraction for ER but a dominant and energy dependent fraction for NR. This complexity distinguishes ER and NR: for the same visible energy, the slower nuclear recoils form a denser track with less charge and more light than recoiling electrons, and they generate fewer long-lived triplet state scintillation photons than singlet-state photons. Charge and light yields depend on drift field, energy, and the initial particle (ER or NR), requiring extensive calibrations. The existing data has been incorporated into the NEST Monte Carlo framework. Typical yields are several tens of electrons and photons per keV of deposited energy (with up to 10-15% efficiency for these photons being detected).

The scattering rates of WIMPs are model dependent, but are usually highest for spin-independent scattering which has an  $A^2$  dependence, so that experiments to date have used LXe and LAr targets. LXe experiments have had the best WIMP sensitivity for most WIMP masses for the last decade, including the current world-leading sensitivity from the 300 kg LUX experiment. Other Xe experiments include XENON10 and XENON100, ZEPLIN III, and PandaX. Next generation experiments under construction include XENON1T with 1 ton fiducial mass, and LZ with 5.6 tons fiducial mass. If a dark matter signal is seen, its spin dependence could be probed with Xe targets isotopically separated into spin-rich and spin-poor targets.

The reach of LXe TPCs depends critically on the level of ER background rejection provided by the ratio of charge to light. Reported values (at 50% NR acceptance) range from 99.6% in LUX to 99.99% in ZEPLIN III, which had a very high (4 kV/cm) drift field. While there is a basic framework [73] for how this improves with light collection and varies with electric field, a fully predictive understanding is not yet in hand. Pulse shape discrimination is present, but weak at low energy. The  $\sim 178 \text{ nm}$  scintillation light of Xe is just long enough to be transmitted through high purity synthetic quartz PMTs windows. Kr suppression to the  $\sim \text{ppt}$  or better level is needed, and has been accomplished via distillation or chromatography.

Two experiments to date have produced dark matter limits using dual phase Ar TPCs: WARP and DarkSide-50, while ArDM is under development. A primary attraction of Ar compared to Xe is much lower raw material costs. However beta decays from  $^{39}\text{Ar}$ , produced by cosmic-ray interactions in the atmosphere, give a low energy ER background roughly  $10^8$  times higher than the fundamental ER background from p-p solar neutrinos. Remarkably, however, pulse shape discrimination (PSD) of ER backgrounds is very powerful in LAr for sufficiently high energy, based on the favorably different ratio of populations of the singlet (6 ns lifetime) and triplet ( $\sim 1.5 \mu\text{s}$  lifetime) states. DarkSide has shown roughly  $10^8$  discrimination with  $\geq 50\%$  WIMP acceptance above 47 keV. They have also extracted “aged” Ar with the 32.9 yr half-life  $^{39}\text{Ar}$  reduced by a factor of 1400, via processing of underground (cosmic ray shielded) gas deposits. This lowers the energy threshold and allows ton-scale experiments without significant pile-up. Charge and light discrimination has also been demonstrated at high energy, but it is less well characterized than in LXe. While the strong PSD in LAr allows relaxed requirements for ER backgrounds, U and Th contamination must still be kept very low because their decay chains create neutrons via  $(\alpha, n)$  reactions, particularly in low  $Z$  elements such as PMT glass and PTFE. Waveshifter is used (typically TPB) because PMTs are blind to the 128 nm scintillation light.

With sufficient control of dissolved Kr and Rn, the ER background

in the next LXe experiments will be the as-yet unmeasured low energy spectrum of solar neutrinos from the main p-p burning reaction. LZ’s sensitivity is about a decade above the “floor” of coherent electron scattering of astrophysical neutrinos, which, absent a directional measurement (see below), are essentially indistinguishable from WIMPs. A 30-50 ton LXe TPC would approach the practical limit set by this floor for WIMP masses above  $\sim 5 \text{ GeV}$  if a  $\sim 99.98\%$  rejection (at 30% NR acceptance) of p-p solar  $\nu$  ER backgrounds [74] is achieved, while a  $\sim 200 \text{ ton}$  LAr detector would achieve similar sensitivity for WIMPs above  $\sim 50 \text{ GeV}$ . Sensitivity to lower WIMP mass could be obtained by adding Ne to a LXe TPC, or, more speculatively, with a superfluid He TPC [75] read out with superconducting sensors (similar to the proposed HERON solar neutrino experiment).

Measurement of NR recoil track direction would provide proof of the galactic origin of a dark matter signal since the prevailing WIMP direction varies on a daily basis as the earth spins. This cannot be achieved for the sub-micron tracks in any existing solid or liquid technology, but the mm-scale tracks in a low pressure gas (typically,  $P \sim 50 \text{ Torr}$ ) could be imaged with sufficiently dense instrumentation. Directionality can be established with  $O(10^2)$  events by measuring just the track direction, while, with finer resolution that distinguishes the diffuse (dense) tail and dense (diffuse) head of NR (ER) tracks, only  $O(10)$  events are required. Such imaging requires a high energy threshold, decreasing WIMP sensitivity, but also powerfully rejecting less dense ER background tracks.

A variety of TPC configurations are being pursued to accomplish this, most with a  $\text{CF}_4$  target. The longest established effort, DRIFT, avoids diffusion washing out tracks for electron drift distances greater than  $\sim 20 \text{ cm}$  by attaching electrons to  $\text{CS}_2$ , which drifts with vastly reduced diffusion. Other efforts drift electrons directly and use a variety of techniques for their measurement: DMTPC (electroluminescence + CCDs), MIMAC (MicroMegas), NEWAGE (GEMs), and  $\text{D}^3$  (Si pixels). WIMP limits have been obtained by DRIFT, NEWAGE, and DMTPC. A related suggestion is that the amount of recombination in a high pressure Xe gas with an electron-cooling additive could be sensitive to the angle between the track and electric field [76], eliminating the need for track imaging. Directional measurements appear to be the only possibility to push beyond the floor of coherent neutrino scatters [77], though this would require very large target mass.

### 35.4.2. $0\nu\beta\beta$ Decay :

Another major class of rare event search is neutrinoless double beta decay ( $0\nu\beta\beta$ ). A limited set of nuclei are unstable against simultaneous beta decay of two neutrons. Observation of the lepton-number violating neutrinoless version of this decay would establish that neutrinos are Majorana particles and provide a direct measure of neutrino mass. For a recent review, see [78]. The signal in  $0\nu\beta\beta$  decay is distinctive: the full Q-value energy of the nuclear decay appears as equal energy back-to-back recoil electrons. A large TPC is advantageous for observing this low rate decay for all the reasons described above. The first detector to observe the standard model process 2 neutrino double beta decay was a gaseous TPC which imaged the two electrons tracks from  $^{82}\text{Se}$  embedded in a foil. Modern detectors use Xe as the detector medium because it includes the  $\beta\beta$  isotope  $^{136}\text{Xe}$  (Q-value 2458 keV), which, as an inert gas, can also be more readily enriched from its natural 8.9% abundance than any other  $\beta\beta$  isotope. EXO-200, which currently has one of the best search limits [79], is a large single-phase LXe TPC with roughly 110 active kg of Xe enriched to 80.7%  $^{136}\text{Xe}$ , and a multi-ton successor nEXO has been proposed which would fully cover the inverted neutrino mass hierarchy. These detectors are similar to dark matter TPCs, but, not needing charge gain, use single phase with charge measured directly on crossed wire grids. Light readout is done with LAAPDs (EXO-200) and SiPMs (nEXO).

The dominant background is gamma rays originating outside the active volume. Most of these undergo multiple Compton-scatters which are efficiently recognized and rejected through sub-cm position resolution, though the few percent of gammas at this energy that photoabsorb are not. Self shielding of gamma rays is less powerful than in dark matter, since in the former case there is some small

probability of penetrating to some depth followed by the modestly small probability of photo-absorption. The latter case consists of three small probability processes: penetration to some depth, a very low-energy scatter, and the gamma exiting without a second interaction. Because of this and the fact that background and the signal are both electron recoils, the requirements on radioactivity in all the materials of a  $\beta\beta$  TPC are much more stringent than an otherwise similar dark matter detector, unless other background rejection tools are available. Percent-level energy resolution is crucial to avoid background from  $2\nu\beta\beta$  decays and gammas including the prominent 2615 MeV line from  $^{208}\text{Tl}$  in the Th chain. Here the combined charge and light measurement eliminates the otherwise dominant fluctuations in recombination which lead to anti-correlated fluctuations in charge and light. EXO-200 has achieved  $\sigma \approx 1.5\%$  (at 2458 keV), and values below 1% appear possible.

The NEXT collaboration uses a high pressure gas phase Xe TPC with electroluminescent readout of the charge to achieve mm spatial resolution so that the two-electron topology of  $0\nu\beta\beta$  events can be distinguished from single electrons from photoabsorption of background gammas. In addition, the low recombination fraction in the gas phase suppresses recombination fluctuations, in principle allowing  $\sigma$  below 0.2% via the charge channel alone. Finally, a definitive identification of a  $0\nu\beta\beta$  signal would be provided by extraction and tagging of the ionized Ba daughter via atomic physics techniques [80], either in gas or liquid phases.

### 35.5. Sub-Kelvin detectors

Written September 2015 by K. Irwin (Stanford and SLAC).

Many particle physics experiments utilize detectors operated at temperatures below 1 K. These include WIMP searches, beta-decay experiments to measure the absolute mass of the electron neutrino, and searches for neutrinoless-double-beta decay ( $0\nu\beta\beta$ ) to probe the properties of Majorana neutrinos. Sub-Kelvin detectors also provide important cosmological constraints on particle physics through sensitive measurement of the cosmic microwave background (CMB). CMB measurements probe the physics of inflation at  $\sim 10^{16}$  GeV, and the absolute mass, hierarchy, and number of neutrino species.

Detectors that operate below 1 K benefit from reduced thermal noise and lower material specific heat and thermal conductivity. At these temperatures, superconducting materials, sensors with high responsivity, and cryogenic preamplifiers and multiplexers are available. We provide a simple overview of the techniques and the experiments using sub-K detectors. A useful review of the broad application of low-temperature detectors is provided in [81], and the proceedings of the International Workshop on Low Temperature Detectors [82] provide an overview of the field.

Sub-Kelvin detectors can be categorized as equilibrium thermal detectors or non-equilibrium detectors. Equilibrium detectors measure a temperature rise in a material when energy is deposited. Non-equilibrium detectors are based on the measurement of prompt, non-equilibrated signals and on the excitation of materials with an energy gap.

#### 35.5.1. Equilibrium thermal detectors :

An equilibrium thermal detector consists of a thermometer and absorber with combined heat capacity  $C$  coupled to a heat bath through a weak thermal conductance  $G$ . The rise time of a thermal detector is limited by the internal equilibration time of the thermometer-absorber system and the electrical time constant of the thermometer. The thermal relaxation time over which heat escapes to the heat bath is  $\tau = C/G$ . Thermal detectors are often designed so that an energy input to the absorber is thermalized and equilibrated through the absorber and thermometer on timescales shorter than  $\tau$ , making the operation particularly simple. An equilibrium thermal detector can be operated as either a calorimeter, which measures an incident energy deposition  $E$ , or as a bolometer, which measures an incident power  $P$ .

In a calorimeter, an energy  $E$  deposited by a particle interaction causes a transient change in the temperature  $\Delta T = E/C$ , where the heat capacity  $C$  can be dominated by the phonons in a lattice, the

quasiparticle excitations in a superconductor, or the electronic heat capacity of a metal. The thermodynamic energy fluctuations in the absorber and thermometer have variance

$$\Delta E_{\text{rms}}^2 = k_{\text{B}} T^2 C \quad (35.5)$$

when operated near equilibrium, where  $\Delta E_{\text{rms}}$  is the root-mean-square energy fluctuation,  $k_{\text{B}}$  is the Boltzmann constant and  $T$  is the equilibrium temperature. When a sufficiently sensitive thermometer is used, and the energy is thermalized at frequencies large compared to the thermal response frequency ( $f_{\text{th}} = 1/2\pi\tau$ ), the signal-to-noise ratio is nonzero at frequencies higher than  $f_{\text{th}}$ . In this case, detector energy resolution can be somewhat better than  $\Delta E_{\text{rms}}$  [83]. Deviations from the ideal calorimeter model can cause excess noise and position and energy dependence in the signal shape, leading to degradation in achieved energy resolution.

In a bolometer, a power  $P$  deposited by a stream of particles causes a change in the equilibrium temperature  $\Delta T = P/G$ . The weak thermal conductance  $G$  to the heat bath is usually limited by the flow of heat through a phonon or electron system. The thermodynamic power fluctuations in the absorber and thermometer have power spectral density

$$S_P = NE P^2 = 4k_{\text{B}} T^2 G \quad (35.6)$$

when operated near equilibrium, where the units of NEP (noise equivalent power) are  $\text{W}/\sqrt{\text{Hz}}$ .

The minimization of thermodynamic energy and power fluctuations is a primary motivation for the use of sub-Kelvin thermal detectors. These low temperatures also enable the use of materials and structures with extremely low  $C$  and  $G$ , and the use of superconducting materials and amplifiers.

When very large absorbers are required (e.g. WIMP dark matter searches), dielectric crystals with extremely low specific heat are often used. These materials are operated well below the Debye temperature  $T_{\text{D}}$  of a crystal, where the specific heat scales as  $T^3$ . In this low-temperature limit, the dimensionless phononic heat capacity at fixed volume reduces to

$$\frac{C_V}{N k_{\text{B}}} = \frac{12\pi^4}{5} \left( \frac{T}{T_{\text{D}}} \right)^3, \quad (35.7)$$

where  $N$  is the number of atoms in the crystal. Normal metals have higher low-temperature specific heat than dielectric crystals, but they also have superior thermalization properties, making them attractive for some applications in which extreme precision and high energy resolution are required (e.g. beta endpoint experiments to measure neutrino mass using  $^{163}\text{Ho}$ ). At low temperature, the heat capacity of normal metals is dominated by electrons, and is linear in temperature, with convenient form

$$C = \frac{\rho}{A} \gamma V T, \quad (35.8)$$

where  $V$  is the sample volume,  $\gamma$  is the molar specific heat of the material,  $\rho$  is the mass density, and  $A$  is the atomic weight. Superconducting absorbers are also used. Superconductors combine some of the thermalization advantages of normal metals with the lower specific heats associated with insulators when operated well below  $T_{\text{c}}$ , where the electronic heat capacity freezes out, and the material is dominated by phononic heat capacity. At higher temperatures, superconducting materials have more complicated heat capacities, but at their transition temperature  $T_{\text{c}}$ , BCS theory predicts that the electronic heat capacity of a superconductor is  $\sim 2.43$  times the normal metal value.

When very low thermal conductances are required for power measurement (e.g. the measurement of the cosmic microwave background), the weak thermal link is sometimes provided by thin membranes of non-stoichiometric silicon nitride. The thermal conductance of these membranes is:

$$G = 4\sigma A T^3 \xi, \quad (35.9)$$

where  $\sigma$  has a value of  $15.7 \text{ mW}/\text{cm}^2 \text{K}^4$ ,  $A$  is the cross-sectional area perpendicular to the heat flow, and  $\xi$  is a numerical factor with a

**Table 35.5:** Some selected experiments using sub-Kelvin equilibrium bolometers to measure the CMB. These experiments constrain the physics of inflation and the absolute mass, hierarchy, and number of neutrino species. The experiment location determines the part of the sky that is observed. The size of the aperture determines the angular resolution. The table also indicates the type of sensor used, the number of sensors, the frequency range, and the number of frequency bands. The number of sensors and frequency range and bands for ongoing upgrades are provided for some experiments in parentheses.

Sub-K CMB Experiment	Location	Aperture	Sensor # type	Sensors (planned)	Frequency (planned)	Bands (planned)
<b>Ground-based</b>						
Atacama Cosmology Telescope (2007–)	Chile	6 m	TES	1,800 (5,334)	90–150 GHz (28–220 GHz)	2 (5)
BICEP/Keck (2006–)	South Pole	26/68 cm	TES	3,200	95–220 GHz	3
CLASS (2015–)	Chile	60 cm	TES	36 (5,108)	40 GHz (40–220 GHz)	1 (4)
POLARBEAR / Simons (2012–)	Chile	3.5 m	TES	1,274 (22,764)	150 GHz (90–220 GHz)	1 (3)
South Pole Telescope (2007–)	South Pole	10 m	TES	1,536 (16,260)	95–150 GHz (95–220 GHz)	2 (3)
<b>Balloon</b>						
EBEX (2013–)	McMurdo	1.5 m	TES	~1,000	150–410 GHz	3
PIPER (2016–)	New Mexico	2 m	TES	5,120	200–600 GHz	4
SPIDER (2014–)	McMurdo	30 cm	TES	1,959	90–280 GHz	3
<b>Satellite</b>						
Planck HFI (2003–)	L2	1.5 m	NTD	52	100–857 GHz	9

value of one in the case of specular surface scattering but less than one for diffuse surface scattering. The thermal impedance between the electron and phonon systems can also limit the thermal conductance.

The most commonly used sub-Kelvin thermometer is the superconducting transition-edge sensor (TES) [84]. The TES consists of a superconductor biased at the transition temperature  $T_c$ , in the region between the superconducting and normal state, where its resistance is a strong function of temperature. The TES is voltage biased. The Joule power provides strong negative electrothermal feedback, which improves linearity, speeds up response to faster than  $\tau = C/G$ , and provides tolerance for  $T_c$  variation between multiple TESs in a large array. The current flowing through a TES is read out by a superconducting quantum interference device (SQUID) amplifier. These amplifiers can be cryogenically multiplexed, allowing a large number of TES devices to be read out with a small number of wires to room temperature.

Neutron-transmutation-doped (NTD) germanium and implanted silicon semiconductors read out by cryogenic FET amplifiers are also used as thermometers [83]. Their electrical resistance is exponentially dependent on  $1/T$ , and is determined by phonon-assisted hopping conduction between impurity sites. Finally, the temperature dependence of the permeability of a paramagnetic material is used as a thermometer. Detectors using these thermometers are referred to as metallic magnetic calorimeters (MMC) [85]. These detectors operate without dissipation and are inductively readout by SQUIDS.

Equilibrium thermal detectors are simple, and they have important advantages in precision measurements because of their insensitivity to statistical variations in energy down-conversion pathways, as long as the incident energy equilibrates into an equilibrium thermal distribution that can be measured by a thermometer.

### 35.5.2. Nonequilibrium Detectors :

Nonequilibrium detectors use many of the same principles and techniques as equilibrium detectors, but are also sensitive to details of the energy down-conversion before thermalization. Sub-Kelvin nonequilibrium detectors measure athermal phonon signals in a dielectric crystal, electron-hole pairs in a semiconductor crystal, athermal quasiparticle excitations in a superconductor, photon emission from a scintillator, or a combination of two of the above to better discriminate recoils from nuclei or electrons. Because the phonons are athermal, sub-Kelvin nonequilibrium detectors can use absorbers with larger heat capacity, and they use information about the details of energy down-conversion pathways in order to better discriminate signal from background.

In WIMP and neutrino experiments using sub-Kelvin dielectric semiconductors, the recoil energy is typically  $\gtrsim 0.1$  keV. The majority of the energy is deposited in phonons and a minority in ionization and, in some cases, scintillation. The semiconductor bandgap is typically  $\sim$  eV, and  $k_B T < 10$   $\mu$ eV at  $T < 1$  K. Thus, high-energy charge pairs and athermal phonons are initially produced. The charge pairs cascade quickly to the gap edge. The high-energy phonons experience isotopic scattering and anharmonic decay, which downshifts the phonon spectrum until the phonon mean free path approaches the characteristic dimension of the absorber. If the crystal is sufficiently pure, these phonons propagate ballistically, preserving information about the interaction location. They are not thermalized, and thus not affected by an increase in the crystal heat capacity, allowing the use of larger absorbers. Sensors similar to those used in sub-K equilibrium thermal detectors measure the athermal phonons at the crystal surface.

Superconductors can also be used as absorbers in sub-Kelvin detectors when  $T \ll T_c$ . The superconducting gap is typically  $\sim$  meV. Energy absorption breaks Cooper pairs and produces quasiparticles. These particles cascade to the superconducting gap edge, and then recombine after a material-dependent lifetime. During the quasiparticle lifetime, they diffuse through the material. In superconductors with large mean free path, the diffusion length can be more than 1 mm, allowing diffusion to a detector.

In some experiments (e.g. SuperCDMS and CRESST), athermal



**Table 35.6:** Selected experiments using sub-Kelvin calorimeters. The table shows only currently operated experiments, and is not exhaustive. WIMP experiments search for dark matter, and beta-decay and neutrinoless double beta decay ( $0\nu\beta\beta$ ) experiments constrain neutrino mass, hierarchy, and Majorana nature. The experiment location determines the characteristics of the radioactive background. The dates of current program phase, detection mode (equilibrium or nonequilibrium phonon measurements, and measurement of ionization or scintillation signals), the absorber and total mass, the sensor type, and the number of sensors and crystals (if different) are given. Many sub-K calorimeter experiments are also in planning and construction phases, including EURECA (dark matter), HOLMES and NuMECs (beta decay), and CUPID-0 ( $0\nu\beta\beta$  decay). Many of the existing experiments are being upgraded to larger mass absorbers, different absorber materials, or lower energy threshold.

Sub-K Calorimeter	Location	Detection mode	Absorber Total mass	Sensor type	# Sensor # Crystal
<b>WIMP</b>					
CRESST II (2003–)	Gran Sasso Italy	Noneq. phon. and scint.	CaWO <sub>4</sub> 5.4 kg	TES	18
EDELWEISS III (2015–)	LSM Modane France	Eq. thermal and ion.	Ge 22 kg	NTD Ge +HEMT	36
SuperCDMS (2012–)	Soudan, USA SNOLAB, Canada	Noneq. phon. and ion.	Ge 9 kg	TES +JFET	120 15
<b>Beta decay</b>					
ECHO (2012–)	Heidelberg Germany	Eq. thermal	Au: <sup>163</sup> Ho 0.2μg	MMC	16
<b><math>0\nu\beta\beta</math> decay</b>					
CUORE (2015–)	Gran Sasso Italy	Eq. thermal	TeO <sub>2</sub> 741 kg	NTD Ge	988
AMoRe Pilot (2015–)	Yang Yang S. Korea	Noneq. phon. and scint.	CaMoO <sub>4</sub> 1.5 kg	MMC	5
LUCIFER (2010–)	Gran Sasso Italy	Eq. thermal and scint.	ZnSe 431 g	NTD Ge	1

phonons and quasiparticle diffusion are combined to increase achievable absorber mass. Athermal phonons in a three-dimensional dielectric crystal break Cooper pairs in a two-dimensional superconducting film on the detector surface. The resulting quasiparticles diffuse to thermal sensors (typically a TES) where they are absorbed and detected. While thin superconducting films have diffusion lengths shorter than the diffusion lengths in single crystal superconductors, segmenting the films into small sections and coupling them to multiple TES sensors allows the instrumentation of large absorber volume. The TES sensors can be wired in parallel to combine their output signal.

The combined measurement of the phonon signal and a secondary signal (ionization or scintillation) can provide a powerful discrimination of signal from background events. Nuclear-recoil events in WIMP searches produce proportionally smaller ionization or scintillation signal than electron-scattering events. Since many of the background events are electron recoils, this discrimination provides a powerful veto. Similarly, beta-decay events produce proportionally smaller scintillation signal than alpha-particle events, allowing rejection of alpha backgrounds in neutrino experiments.

Combined phonon and ionization measurement has been implemented in experiments including CDMS I/II, SuperCDMS, and EDELWEISS I/II/III. These experiments use semiconductor crystal absorbers, in which dark-matter scattering events would produce recoiling particles and generate electron-hole pairs and phonons. The electron-hole pairs are separated and drifted to the surface of the crystal by applying an electric field, where they are measured by a JFET or HEMT using similar techniques to those used in 77 K Ge x-ray spectrometers. However, the field strength must be much lower in sub-K detectors to limit the generation of phonon signals by the Neganov-Luke effect, which can confuse the background discrimination. For detectors with very low threshold, the Neganov-Luke effect can also be used to detect generated charge through the induced phonon signal.

Combined phonon and scintillation measurement has been implemented in CRESST II, ROSEBUD, AMoRE and LUCIFER. For example, the CRESST-II experiment uses CaWO<sub>4</sub> crystal absorbers, and measures both the phonon signal and the scintillation signal with TES calorimeters. A wide variety of scintillating crystals are under consideration, including different tungstates and molybdates, BaF<sub>2</sub>, ZnSe, and bismuth germanate (BGO).

### 35.6. Low-radioactivity background techniques

Revised August 2015 by A. Piepke (University of Alabama).

The physics reach of low-energy rare-event searches *e.g.* for dark matter, neutrino oscillations, or double beta decay is often limited by background caused by radioactivity. Depending on the chosen detector design, the separation of the physics signal from this unwanted interference can be achieved on an event-by-event basis by active event tagging, utilizing some unique event features, or by reducing the flux of the background radiation by appropriate shielding and material selection. In both cases, the background rate is proportional to the flux of the interfering radiation. Its reduction is thus essential for realizing the full physics potential of the experiment. In this context, “low energy” may be defined as the regime of natural, anthropogenic, or cosmogenic radioactivity, all at energies up to about 10 MeV. See [86,87] for in-depth reviews of this subject. Following the classification of [86], sources of background may be categorized into the following classes:

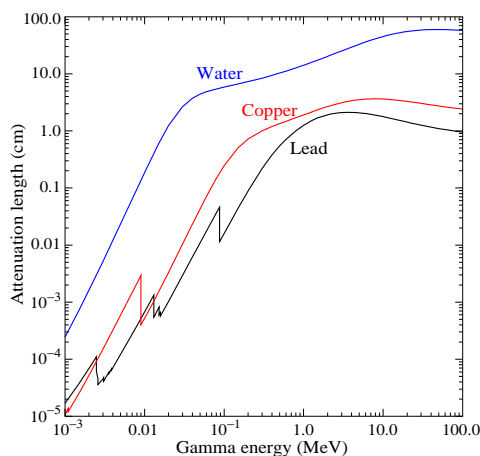
1. environmental radioactivity,
2. radio-impurities in detector or shielding components,
3. radon and its progeny,
4. cosmic rays,
5. neutrons from natural fission, ( $\alpha, n$ ) reactions and from cosmic-ray muon spallation and capture.

**35.6.1. Defining the problem :** The application defines the requirements. Background goals can be as demanding as a few low-energy events per year in a ton-size detector. The strength of the physics signal of interest can often be estimated theoretically or from limits derived by earlier experiments. The experiments are then designed for the desired signal-to-background ratio. This requires finding the right balance between “clarity of measurement”, ease of construction, and budget. In a practical sense, it is important to formulate background goals that are sufficient for the task at hand but achievable, in a finite time. It is standard practice to use detector simulations to translate the background requirements into limits for the radioactivity content of various detector components, requirements for radiation shielding, and allowable cosmic-ray fluxes. This strategy allows the identification of the most critical components early and facilitates the allocation of analysis and development resources in a rational way. The CERN code GEANT4 [88] is a widely used tool for this purpose. It has incorporated sufficient nuclear physics to allow accurate background estimations. Custom-written event generators, modeling *e.g.*, particle correlations in complex decay schemes, deviations from allowed beta spectra or  $\gamma$ - $\gamma$ -angular correlations, are used as well.

**35.6.2. Environmental radioactivity :** The long-lived natural radio-nuclides  $^{40}\text{K}$ ,  $^{232}\text{Th}$ , and  $^{238}\text{U}$  have average abundances of 1.6, 11.1 and 2.7 ppm (corresponding to 412, 45 and 33 Bq/kg, respectively) in the earth’s crust, with large local variations. In most applications,  $\gamma$  radiation emitted due to the decay of natural radioactivity and its unstable daughters constitutes the dominant contribution to the local radiation field. Typical low-background applications require levels of natural radioactivity on the order of ppb or ppt in the detector components. Passive or active shielding is used to suppress external  $\gamma$  radiation down to an equivalent level. Fig. 35.9 shows the energy-dependent attenuation length  $\lambda(E_\gamma)$  as a function of  $\gamma$ -ray energy  $E_\gamma$  for three common shielding materials (water, copper, lead). The thickness  $\ell$  required to reduce the external flux by a factor  $f > 1$  is estimated, assuming exponential damping:

$$\ell = \lambda(E_\gamma) \cdot \ln f. \quad (35.10)$$

At 100 keV, a typical energy scale for dark matter searches (or 2.615 MeV, for a typical double-beta decay experiment), attenuation by a factor  $f = 10^5$  requires 67(269) cm of  $\text{H}_2\text{O}$ , 2.8(34) cm of Cu, or 0.18(23) cm of Pb. Such estimates allow for an order-of-magnitude determination of the experiment dimensions.



**Figure 35.9:**  $\gamma$ -ray attenuation lengths in some common shielding materials. The mass attenuation data has been taken from the NIST data base XCOM; see “Atomic Nuclear Properties” at [pdg.lbl.gov](http://pdg.lbl.gov).

A precise estimation of the the magnitude of the external gamma-ray background, including scattering and the effect of analysis-energy cuts, requires Monte Carlo simulations based on the the knowledge

of the radioactivity present in the laboratory. Detailed modeling of the  $\gamma$ -ray flux in a large laboratory, or inside the hermetic shielding, needs to cope with a very small probability of generating any signal in the detector. It is often advantageous to calculate solid angle of the detector to the background sources and mass attenuation of the radiation shield separately, or to employ importance sampling. The former method can lead to loss of energy-direction correlations while in the latter has to balance CPU-time consumption against the loss of statistical independence. These approaches reduce the computation time required for a statistically meaningful number of detector hits to manageable levels.

Water is commonly used as shielding medium for large detectors, as it can be obtained cheaply and purified effectively in large quantity. Water purification technology is commercially available. Ultra-pure water, instrumented with photomultiplier tubes, can serve as active cosmic-ray veto counter. Water is also an effective neutron moderator and shield. In more recent underground experiments that involve detectors operating at cryogenic temperature, liquefied gases (*e.g.* argon) are being used for shielding as well.

**35.6.3. Radioactive impurities in detector and shielding components :** After suppressing the effect of external radioactivity, radioactive impurities, contained in the detector components or attached to their surfaces, become important. Every material contains radioactivity at some level. The activity can be natural, cosmogenic, man-made, or a combination of them. The determination of the activity content of a specific material or component requires case-by-case analyses, and is rarely obtainable from the manufacturer. However, there are some general rules that can be used to guide the pre-selection. For detectors designed to look for electrons (for example in double-beta decay searches or neutrino detection via inverse beta decay or elastic scattering), intrinsic radioactivity is often the principal source of background. For devices detecting nuclear recoils (for example in dark matter searches), this is often of secondary importance as ionization signals can be actively discriminated on an event-by-event basis. Decay induced nuclear reactions become a concern.

For natural radioactivity, a rule of thumb is that synthetic substances are cleaner than natural materials. Typically, more highly processed materials have lower activity content than raw substances. Substances with high electro-negativity tend to be cleaner as the refining process preferentially removes K, Th, and U. For example, Al is often found to contain considerable amounts of Th and U, while electrolytic Cu is very low in primordial activities. Plastics or liquid hydrocarbons, having been refined by distillation, are often quite radiopure. Tabulated radioassay results for a wide range of materials can be found in Refs. [89] and [90]. Radioassay results from previous underground physics experiments are being archived at an online database [91].

The long-lived  $^{238}\text{U}$  daughter  $^{210}\text{Pb}$  ( $T_{1/2}=22.3$  y) is found in all shielding lead, and is a background concern at low energies. This is due to the relatively high endpoint energy ( $Q_\beta=1.162$  MeV) of its beta-unstable daughter  $^{210}\text{Bi}$ . Lead refined from selected low-U ores have specific activities of about 5–30 Bq/kg. For applications that require lower specific activity, ancient lead (for example from Roman ships) is sometimes used. Because the ore processing and lead refining removed most of the  $^{238}\text{U}$ , the  $^{210}\text{Pb}$  decayed during the long waiting time to the level supported by the U-content of the refined lead. Lining the lead with copper to range out the low-energy radiation is another remedy. However, intermediate- $Z$  materials carry additional cosmogenic-activation risks when handled above ground, as will be discussed below.  $^{210}\text{Pb}$  is also found in solders.

Man-made radioactivity, released during above-ground nuclear testing and nuclear power production, is a source of background. The fission product  $^{137}\text{Cs}$  can often be found attached to the surface of materials. The radioactive noble gas  $^{85}\text{Kr}$ , released into the atmosphere by nuclear reactors and nuclear fuel re-processing, is sometimes a background concern, especially due to its high solubility in organic materials. Post-World War II steel typically contains a few tens of mBq/kg of  $^{60}\text{Co}$ .

Surface activity is not a material property per se but is added

during manufacturing and handling. Surface contamination can often be effectively removed by clean machining, etching, or a combination of both. The assembly of low-background detectors is often performed in controlled enclosures (e.g. clean rooms or glove boxes) to avoid contaminating surfaces with environmental substances, such as dust, containing radioactivity at much higher concentrations than the detector components. Surfaces are cleaned with high purity chemicals and de-ionized water. When not being processed components are best stored in sealed bags to limit dust deposition on the surface, even inside clean rooms. Surface contamination can be quantified by means of wipe-testing with acid or alcohol wetted Whatman 41 filters. Pre-soaking of the filters in clean acid reduces the amount of Th and U contained in the paper and boosts analysis sensitivity. The paper filters are ashed after wiping and the residue is digested in acid. Subsequent analysis by means of mass spectroscopy or neutron activation analysis is capable of detecting less than 1 pg/cm<sup>2</sup> of Th and U.

The most demanding low-rate experiments require screening of *all* components, which can be a time consuming task. The requirements for activity characterization depend on the experiment and the location and amount of a particular component. Monte Carlo simulations are used to quantify these requirements. Sensitivities of the order  $\mu\text{Bq/kg}$  or less are sometimes required for the most critical detector components. At such a level of sensitivity, the characterization becomes a challenging problem in itself. Low-background  $\alpha$ ,  $\beta$ , and  $\gamma$ -ray counting, mass spectroscopy, and neutron activation analysis are the commonly used diagnostic techniques.

**35.6.4. Radon and its progeny :** The noble gas <sup>222</sup>Rn, a pure  $\alpha$ -emitter, is a <sup>238</sup>U decay product. Due to its relatively long half-life of 3.8 d it is released by surface soil and is found in the atmosphere everywhere. <sup>220</sup>Rn (<sup>232</sup>Th decay product) is mostly unimportant for most low-background experiments because of its short half-life. The <sup>222</sup>Rn activity in air ranges from 10 to 100 mBq/L outdoors and 100 to thousands of mBq/L indoors. The natural radon concentration depends on the weather and shows daily and seasonal variations. Radon levels are lowest above the oceans. For electron detectors, it is not the Rn itself that creates background, but its progeny <sup>214</sup>Pb, <sup>214</sup>Bi, <sup>210</sup>Bi, which emit energetic beta and  $\gamma$  radiation. Thus, not only the detector itself has to be separated from contact with air, but also internal voids in the shield which contain air can be a background concern. Radon is quite soluble in water and even more so in organic solvents. For large liquid scintillation detectors, radon mobility due to convection and diffusion is a concern. To define a scale: typical double-beta-decay searches are restricted to  $< \mu\text{Bq/kg}_{\text{detector}}$  (or 1 decay per kg<sub>detector</sub> and per 11.6 days) activities of <sup>222</sup>Rn in the active medium. This corresponds to a steady-state population of 0.5 atoms/kg<sub>detector</sub> or 50  $\mu\text{L/kg}_{\text{detector}}$  of air (assuming 20 mBq/L of radon in the air). The demand on leak tightness can thus be quite demanding. The decay of Rn itself is a concern for some recoil type detectors, as nuclear recoil energies in  $\alpha$  decays are substantial (76 keV in the case of <sup>222</sup>Rn).

Low-background detectors are often kept sealed from the air and continuously flushed with boil-off nitrogen, which contains only small amounts of Rn. For the most demanding applications, the nitrogen is purified by multiple distillations, or by using pressure swing adsorption chromatography. Then only the Rn outgassing of the piping (due to its intrinsic U content) determines the radon concentration. Radon diffuses readily through thin plastic barriers. If the detector is to be isolated from its environment by means of a membrane, the choice of material is important [92].

Prolonged exposure of detector components or raw materials to air leads to the accumulation of the long-lived radon daughter <sup>210</sup>Pb on surfaces. Due to its low Q-value of 63.5 keV, <sup>210</sup>Pb itself is only a problem when extreme low energy response is important. However, because of its higher Q-value, the lead daughter <sup>210</sup>Bi, is a concern up to the MeV scale. The alpha unstable Bi-daughter <sup>210</sup>Po ( $E_\alpha = 5304$  keV) contributes not only to the alpha background but can also induce the emission of energetic neutrons via  $(\alpha, n)$  reactions on low-Z materials (such as F, C, Si...etc). The neutrons, in turn, may capture on other detector components, creating energetic background.

The  $(\alpha, n)$  reaction yield induced by the  $\alpha$  decay of <sup>210</sup>Po is typically small ( $6 \cdot 10^{-6} n/\alpha$  in Teflon, for example). Some data is available on the deposition of radon daughters from air onto materials, see e.g. [94]. This data indicates effective radon daughter collection distances of a few cm in air. These considerations limit the allowable air exposure time. In case raw materials (e.g. in the form of granules) were exposed to air at the production site, the bulk of the finished detector components may be loaded with <sup>210</sup>Pb and its daughters. These are difficult to detect as no energetic gamma radiation is emitted in their decays. Careful air-exposure management is the only way to reduce this source of background. This can be achieved by storing the parts under a protective low-radon cover gas or keeping them sealed from radon.

State-of-the-art detectors can detect radon even at the level of few atoms. Solid state, scintillation, or gas detectors utilize alpha spectroscopy or are exploiting the fast  $\beta - \alpha$  decay sequences of <sup>214</sup>Bi and <sup>214</sup>Po. The efficiency of these devices is sometimes boosted by electrostatic collection of charged radon from a large gas volume into a small detector.

**35.6.5. Cosmic rays :** Cosmic radiation, discussed in detail in Chapter 29, is a source of background for just about any non-accelerator experiment. Primary cosmic rays are about 90% protons, 9% alpha particles, and the rest heavier nuclei (Fig. 29.1). They are totally attenuated within the first few hg/cm<sup>2</sup> of atmospheric thickness. At sea level secondary particles ( $\pi^\pm : p : e^\pm : n : \mu^\pm$ ) are observed with relative intensities 1 : 13 : 340 : 480 : 1420 (Ref. 95; also see Fig. 29.4).

All but the muon and the neutron components are readily absorbed by overburden such as building ceilings and passive shielding. Only if there is very little overburden ( $\lesssim 10$  g/cm<sup>2</sup> or so [86]) do pions and protons need to be considered when estimating the production rate of cosmogenic radioactivity.

Sensitive experiments are thus operated deep underground where essentially only muons can penetrate. As shown in Fig. 29.7, the muon intensity falls off rapidly with depth. Active detection systems, capable of tagging events correlated in time with cosmic-ray activity, are needed, depending on the overburden.

The muonic background is related to low-radioactivity techniques insofar as photo-nuclear interactions with atomic nuclei can produce long-lived radioactivity directly or indirectly via the creation of neutrons. This happens at any overburden, however, at strongly depth dependent rates. Muon bremsstrahlung, created in high-Z shielding materials, contributes to the low energy background too. Active muon detection systems are effective in reducing this background, but only for activities with sufficiently short half-lives, allowing vetoing with reasonable detector dead time.

Cosmogenic activation of detector components at the surface can be an issue for low-background experiments. Proper management of parts and materials above ground during manufacturing and detector assembly minimizes the accumulation of long-lived activity. Cosmogenic activation is most important for intermediate-Z materials such as Cu and Fe. For the most demanding applications, metals are stored and transported under sufficient shielding to stop the hadronic component of the cosmic rays. Parts can be stored underground for long periods before being used. Underground machine shops are sometimes used to limit the duration of exposure at the surface. Some experiments are even electro-forming copper underground.

**35.6.6. Neutrons :** Neutrons contribute to the background of low-energy experiments in different ways: directly through nuclear recoil in the detector medium, and indirectly, through the production of radio-nuclides, capture  $\gamma$ s and inelastically scattering inside the detector and its components. The indirect mechanisms allow even remote materials to contribute to the background by means of penetrating  $\gamma$  radiation. Neutrons are thus an important source of low-energy background. They are produced in different ways:

1. At the earth's surface the flux of cosmic-ray secondary neutrons is exceeded only by that of muons;
2. Energetic tertiary neutrons are produced by cosmic-ray muons by nuclear spallation in the detector and laboratory walls;

3. In high- $Z$  materials, often used in radiation shields, nuclear capture of negative muons results in the emission of neutrons;
4. Natural radioactivity has a neutron component through spontaneous fission and  $(\alpha, n)$ -reactions.

A calculation with the hadronic simulation code FLUKA [93], using the known energy distribution of secondary neutrons at the earth's surface [96], yields a mass attenuation of  $1.5 \text{ hg/cm}^2$  in concrete for secondary neutrons. In case energy-dependent neutron-capture cross sections are known, such calculations can be used to obtain the production rate of particular radio-nuclides.

At an overburden of only few meters water equivalent, neutron production by muons becomes the dominant mechanism. Neutron production rates are high in high- $Z$  shielding materials. A high- $Z$  radiation shield, discussed earlier as being effective in reducing background due to external radioactivity, thus acts as a source for cosmogenic tertiary high-energy neutrons. Depending on the overburden and the radioactivity content of the laboratory, there is an optimal shielding thickness. Water shields, although bulky, are an attractive alternative due to their low neutron production yield and self-shielding.

Shields made from plastic or water are commonly used to reduce the neutron flux. The shield is sometimes doped with a substance having a high thermal neutron capture cross section (such as boron) to absorb thermal neutrons more quickly. The hydrogen, contained in these shields, serves as a target for elastic scattering, and is effective in reducing the neutron energy. Neutrons from natural radioactivity have relatively low energies and can be effectively suppressed by a neutron shield. Ideally, such a neutron shield should be inside the lead to be effective for tertiary neutrons. However, this is rarely done as it increases the neutron production target (in form of the passive shield), and the costs increase as the cube of the linear dimensions. An active cosmic-ray veto is an effective solution, correlating a neutron with its parent muon. This solution works best if the veto system is as far away from the detector as feasible (outside the radiation shield) in order to correlate as many background-producing muons with neutrons as possible. The vetoed time after a muon hit needs to be sufficiently long to assure muon bremsstrahlung and neutron-induced backgrounds are sufficiently suppressed. An upper limit to the allowable veto period is given by the veto-induced deadtime, which is related to the muon hit rate on the veto detector. This consideration also constitutes the limiting factor for the physical size of the veto system (besides the cost). The background caused by neutron-induced radioactivity with live-times exceeding the veto time cannot be addressed in this way. Moving the detector deep underground, and thus reducing the muon flux, is the only technique that addresses all sources of cosmogenic the neutron background.

#### References:

1. R.M. Baltrusaitis *et al.*, Nucl. Instrum. Methods **A20**, 410 (1985).
2. T. Abu-Zayyad *et al.*, Nucl. Instrum. Methods **A450**, 253 (2000).
3. H. Tokuno *et al.*, Nucl. Instrum. Methods **A676**, 54 (2012).
4. J. Abraham *et al.* [Pierre Auger Collab.], Nucl. Instrum. Methods **A620**, 227 (2010).
5. J. Abraham *et al.* [Pierre Auger Collab.], Eur. Phys. J. Plus **127**, 94 (2012).
6. A. Huang, G. Medina-Tanco and A. Santangelo, Experimental Astronomy **40**, 1 (2015).
7. F. Arqueros, J. Hrandel, and B. Keilhauer, Nucl. Instrum. Methods **A597**, 23 (2008).
8. F. Arqueros, J. Hrandel, and B. Keilhauer, Nucl. Instrum. Methods **A597**, 1 (2008).
9. J. Rosado, F. Blanco, and F. Arqueros, Astropart. Phys. **34**, 164 (2010).
10. J. Boyer *et al.*, Nucl. Instrum. Methods **A482**, 457 (2002); J. Abraham *et al.* [Pierre Auger Collab.], Nucl. Instrum. Methods **A620**, 227 (2010).
11. M. Ave *et al.* [AIRFLY Collab.], Astropart. Phys. **28**, 41 (2007).
12. J. Rosado, F. Blanci, and F. Arqueros, Astropart. Phys. **55**, 51 (2014).
13. J.T. Brack *et al.*, Astropart. Phys. **20**, 653, (2004).
14. B. Fick *et al.*, JINST **1**, 11003 (2006).
15. J. Abraham *et al.* [Pierre Auger Collab.], Astropart. Phys. **33**, 108 (2010).
16. J. Abraham *et al.* [Pierre Auger Collab.], Astropart. Phys. **34**, 368 (2011).
17. M. Unger *et al.*, Nucl. Instrum. Methods **A588**, 433 (2008).
18. T.K. Gaisser and A.M. Hillas, Proc. 15th Int. Cosmic Ray Conf. Bulgarska Akademiia na Naukite, Conf. Papers **8**, 353 (1978).
19. J. Holder *et al.*, AIP Conf. Proc. **1085**, 657 (2008).
20. F. Aharonian *et al.*, Astron. & Astrophys. **457**, 899 (2006).
21. J. Albert *et al.*, Astrophys. J. **674**, 1037 (2008).
22. F. Aharonian *et al.*, Astron. & Astrophys. **464**, 235 (2007).
23. E. Aliu *et al.*, Science **322**, 1221 (2008).
24. T.C. Weekes *et al.*, Astrophys. J. **342**, 379 (1989).
25. A.M. Hillas *et al.*, Astrophys. J. **503**, 744 (1998).
26. <http://tevcat.uchicago.edu/>.
27. F.A. Aharonian *et al.*, Astrophys. J. **636**, 777 (2006).
28. B.S. Acharya *et al.*, Astropart. Phys. **43**, 3 (2013).
29. L.A. Bernstein *et al.*, arXiv:0907.4183 (2009).
30. Y. Ashie *et al.*, Phys. Rev. **D71**, 112005 (2005).
31. S. Kasuga *et al.*, Phys. Lett. **B374**, 238 (1996).
32. M. Shiozawa, Nucl. Instrum. Methods **A433**, 240 (1999).
33. K. Abe *et al.*, Phys. Rev. **D83**, 052010 (2011).
34. J. Beacom and M. Vagins, Phys. Rev. Lett. **93**, 171101 (2004).
35. J. Boger *et al.*, Nucl. Instrum. Methods **A449**, 172 (2000).
36. T.K. Gaisser, F. Halzen, and T. Stanev, Phys. Reports **258**, 173 (1995); T.K. Gaisser, F. Halzen, and T. Stanev, Phys. Reports **271**, 355 (1995).
37. J.G. Learned and K. Mannheim, Ann. Rev. Nucl. and Part. Sci. **50**, 679 (2000).
38. U.F. Katz and C. Spiering, Prog. in Part. Nucl. Phys. **67**, 651 (2012).
39. A. Avronin *et al.* [Baikal Collab.], Phys. Part. Nucl. **46**, 211 (2015).
40. M.G. Aartsen *et al.* [IceCube-PINGU Collab.], arXiv:1401.2046 (2014).
41. M.G. Aartsen *et al.* [IceCube-Gen2 Collab.], arXiv:1412.5106 (2014).
42. S. Adrián-Martínez *et al.* [KM3NeT Collab.], Eur. Phys. J. **C74**, 3056 (2014); S. Adrián-Martínez *et al.* [KM3NeT Collab.], Eur. Phys. J. **C76**, 54 (2016).
43. M.G. Aartsen *et al.* [IceCube Collab.], Phys. Rev. Lett. **113**, 101101 (2014).
44. S. Adrián-Martínez *et al.* [ANTARES Collab.], Astrophys. J. **786**, L5 (2014).
45. M.G. Aartsen *et al.* [IceCube Collab.], Phys. Rev. **D91**, 072004 (2015).
46. M.G. Aartsen *et al.* [IceCube Collab.], arXiv:1510.05222-05223, arXiv:1510.05225-05227 (2015).
47. C.W. James, PoS(ICRC2015)024 (2015).
48. K. Griesen, Phys. Rev. Lett. **16**, 748 (1966); G.T. Zatsepin and V.A. Kuzmin, JETP Lett. **4**, 78 (1966).
49. R. Abbasi *et al.* [IceCube Collab.], Astropart. Phys. **34**, 382 (2011).
50. S.R. Klein, arXiv:1012.1407 (2010).
51. G.A. Askaryan, Sov. Phys. JETP **14**, 441 (1962); G.A. Askaryan, Sov. Phys. JETP **21**, 658 (1965).
52. J. Alvarez-Muniz *et al.*, Phys. Rev. **D62**, 063001 (2000).
53. D. Saltzberg *et al.*, Phys. Rev. Lett. **86**, 2802 (2001); O. Scholten *et al.*, J. Phys. Conf. Ser. **81**, 012004 (2007).
54. P. Allison *et al.* [ARA Collab.], Astropart. Phys. **70**, 62 (2015).
55. J.C. Hanson *et al.* [ARIANNA Collab.], J. Glaciology **61**, 438 (2015).
56. J. Avva *et al.*, arXiv:1409.5413 (2014).
57. T. Huege, Braz. J. Phys. **44**, 520 (2014).
58. S. Hoover *et al.*, Phys. Rev. Lett. **105**, 151101 (2010).
59. M. Detrixhe *et al.* [ANITA-II Collab.], Phys. Rev. **D83**, 023513 (2011).

60. R.D. Dagkesamanskii and I.M. Zheleznykh, Sov. Phys. JETP Lett. **50**, 233 (1989).
61. J.D. Bray *et al.*, [arXiv:1509.05256](#) (2015).
62. J.A. Dowdeswell and S. Evans, Rept. on Prog. in Phys. **67**, 1821 (2004).
63. P. Gorham *et al.* [ANITA Collab.], Phys. Rev. Lett. **103**, 051103 (2009). The published limit is corrected in an erratum, P. Gorham *et al.*, [arXiv:1011.5004](#) (2010).
64. I. Kravchenko *et al.*, Phys. Rev. **D73**, 082002 (2006); I. Kravchenko *et al.*, Astropart. Phys. **19**, 15 (2003).
65. P. Allison *et al.* [ARA Collab.], [arXiv:1507.08991](#) (2015); P. Allison *et al.* [ARA Collab.], Astropart. Phys. **70**, 62 (2015).
66. S.W. Barwick *et al.* [ARIANNA Collab.], [arXiv:1410.7369](#) (2014).
67. S.A. Wissel *et al.*, PoS(ICRC2015)1150 (2015).
68. J.D. Bray *et al.*, Phys. Rev. **D91**, 063002 (2015).
69. S.W. Barwick *et al.* [ARIANNA Collab.], Astropart. Phys. **70**, 12 (2015).
70. E. Aprile and T. Doke, Rev. Mod. Phys. **82**, 2053 (2010).
71. V. Chepel and H. Araújo, J. Instrum. **8**, R04001 (2013).
72. T.M. Undagoitia and L. Rauch, [arXiv:1509.08767](#) (2015).
73. C.E. Dahl, PhD thesis, Princeton U. (2009).
74. M. Schumann *et al.*, JCAP **10**, 16 (2015).
75. W. Guo and D.N. McKinsey, Phys. Rev. **D87**, 115001 (2013).
76. D.R. Nygren, J. Phys. Conf. Ser. **460**, 012006 (2013).
77. C.A.J. O'Hare *et al.*, Phys. Rev. **D92**, 063518 (2015).
78. S.M. Bilenky and C. Giunti, Mod. Phys. Lett. **A27**, 1230015 (2012).
79. J.B. Albert *et al.*, Nature **510**, 229 (2014).
80. M.L. Moe, Phys. Rev. **C44**, 931 (1991).
81. C. Enss (ed.), *Cryogenic Particle Detection*, Springer-Verlag, Berlin (2005).
82. E. Shirokoff (ed.), *Proc. 15th Int. Workshop on Low Temperature Detectors (LTD-15)*, J. Low Temp. Phys. **176**, 131–1108 (2014); see also previous *Proceedings* of this workshop.
83. S.H. Moseley, J.C. Mather, and D. McCammon, J. Appl. Phys. **56**, 1257 (1984).
84. K.D. Irwin, Appl. Phys. Lett. **66**, 1998 (1995).
85. S.R. Bandler *et al.*, J. Low. Temp. Phys. **93**, 709 (1993).
86. G. Heusser, Ann. Rev. Nucl. and Part. Sci. **45**, 543 (1995).
87. J.A. Formaggio and C.J. Martoff, Ann. Rev. Nucl. and Part. Sci. **54**, 361 (2004).
88. S. Agostinelli *et al.*, Nucl. Instrum. Methods **A506**, 250 (2003).
89. P. Jagam and J.J. Simpson, Nucl. Instrum. Methods **A324**, 389 (1993).
90. D.S. Leonard *et al.*, Nucl. Instrum. Methods **A591**, 490 (2008).
91. <http://www.radiopurity.org>.
92. M. Wojcik *et al.*, Nucl. Instrum. Methods **A449**, 158 (2000).
93. <http://www.fluka.org/fluka.php?id=faq&sub=13>.
94. V.E. Guiseppe *et al.*, [arXiv:1101.0126](#) (2011).
95. National Council on Radiation Protection and Measurement, Report 94, Bethesda, MD (1987).
96. M.S. Gordon *et al.*, IEEE Trans. **NS51**, 3427 (2004).

## 36. RADIOACTIVITY AND RADIATION PROTECTION

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### 36.1. Definitions [1,2]

It would be desirable if legal protection limits could be expressed in directly measurable *physical quantities*. However, this does not allow to quantify biological effects of the exposure of the human body to ionizing radiation.

For this reason, protection limits are expressed in terms of so-called *protection quantities* which, although calculable, are not measurable. Protection quantities quantify the extent of exposure of the human body to ionizing radiation from both whole and partial body external irradiation and from intakes of radionuclides.

In order to demonstrate compliance with dose limits, so-called *operational quantities* are typically used which aim at providing conservative estimates of protection quantities. Often radiation protection detectors used for individual and area monitoring are calibrated in terms of operational quantities and, thus, these quantities become “measurable”.

#### 36.1.1. Physical quantities :

• **Fluence,  $\Phi$**  (unit:  $1/\text{m}^2$ ): The fluence is the quotient of  $dN$  by  $da$ , where  $dN$  is the number of particles incident upon a small sphere of cross-sectional area  $da$

$$\Phi = dN/da . \quad (36.1)$$

In dosimetric calculations, fluence is frequently expressed in terms of the lengths of the particle trajectories. It can be shown that the fluence,  $\Phi$ , is given by

$$\Phi = dl/dV ,$$

where  $dl$  is the sum of the particle trajectory lengths in the volume  $dV$ .

• **Absorbed dose,  $D$**  (unit: gray,  $1 \text{ Gy} = 1 \text{ J/kg} = 100 \text{ rad}$ ): The absorbed dose is the energy imparted by ionizing radiation in a volume element of a specified material divided by the mass of this volume element.

• **Kerma,  $K$**  (unit: gray): Kerma is the sum of the initial kinetic energies of all charged particles liberated by indirectly ionizing radiation in a volume element of the specified material divided by the mass of this volume element.

• **Linear energy transfer,  $L$  or  $LET$**  (unit:  $\text{J/m}$ , often given in  $\text{keV}/\mu\text{m}$ ,  $1 \text{ keV}/\mu\text{m} \approx 1.602 \times 10^{-10} \text{ J/m}$ ): The linear energy transfer is the mean energy,  $dE$ , lost by a charged particle owing to collisions with electrons in traversing a distance  $dl$  in matter. *Low-LET radiation*: X rays and gamma rays (accompanied by charged particles due to interactions with the surrounding medium) or light charged particles such as electrons that produce sparse ionizing events far apart at a molecular scale ( $L < 10 \text{ keV}/\mu\text{m}$ ). *High-LET radiation*: neutrons and heavy charged particles that produce ionizing events densely spaced at a molecular scale ( $L > 10 \text{ keV}/\mu\text{m}$ ).

• **Activity,  $A$**  (unit: becquerel,  $1 \text{ Bq} = 1/\text{s} = 27 \text{ pCi}$ ): Activity is the expectation value of the number of nuclear decays occurring in a given quantity of material per unit time.

#### 36.1.2. Protection quantities :

• **Organ absorbed dose,  $D_T$**  (unit: gray): The mean absorbed dose in an organ or tissue  $T$  of mass  $m_T$  is defined as

$$D_T = \frac{1}{m_T} \int_{m_T} D dm .$$

• **Equivalent dose,  $H_T$**  (unit: sievert,  $1 \text{ Sv} = 100 \text{ rem}$ ): The equivalent dose  $H_T$  in an organ or tissue  $T$  is equal to the sum of the absorbed doses  $D_{T,R}$  in the organ or tissue caused by different radiation types  $R$  weighted with so-called radiation weighting factors  $w_R$ :

$$H_T = \sum_R w_R \times D_{T,R} . \quad (36.2)$$

**Table 36.1:** Radiation weighting factors,  $w_R$ .

Radiation type	$w_R$
Photons, electrons and muons	1
Neutrons, $E_n < 1 \text{ MeV}$	$2.5 + 18.2 \times \exp[-(\ln E_n)^2/6]$
$1 \text{ MeV} \leq E_n \leq 50 \text{ MeV}$	$5.0 + 17.0 \times \exp[-(\ln(2E_n))^2/6]$
$E_n > 50 \text{ MeV}$	$2.5 + 3.25 \times \exp[-(\ln(0.04E_n))^2/6]$
Protons and charged pions	2
Alpha particles, fission fragments, heavy ions	20

It expresses long-term risks (primarily cancer and leukemia) from low-level chronic exposure. The values for  $w_R$  recommended by ICRP [2] are given in Table 36.1.

• **Effective dose,  $E$**  (unit: sievert): The sum of the equivalent doses, weighted by the tissue weighting factors  $w_T$  ( $\sum_T w_T = 1$ ) of several organs and tissues  $T$  of the body that are considered to be most sensitive [2], is called “effective dose”:

$$E = \sum_T w_T \times H_T . \quad (36.3)$$

#### 36.1.3. Operational quantities :

• **Ambient dose equivalent,  $H^*(10)$**  (unit: sievert): The dose equivalent at a point in a radiation field that would be produced by the corresponding expanded and aligned field in a 30 cm diameter sphere of unit density tissue (ICRU sphere) at a depth of 10 mm on the radius vector opposing the direction of the aligned field. Ambient dose equivalent is the operational quantity for *area monitoring*.

• **Personal dose equivalent,  $H_p(d)$**  (unit: sievert): The dose equivalent in ICRU tissue at an appropriate depth,  $d$ , below a specified point on the human body. The specified point is normally taken to be where the individual dosimeter is worn. For the assessment of effective dose,  $H_p(10)$  with a depth  $d = 10 \text{ mm}$  is chosen, and for the assessment of the dose to the skin and to the hands and feet the personal dose equivalent,  $H_p(0.07)$ , with a depth  $d = 0.07 \text{ mm}$ , is used. Personal dose equivalent is the operational quantity for *individual monitoring*.

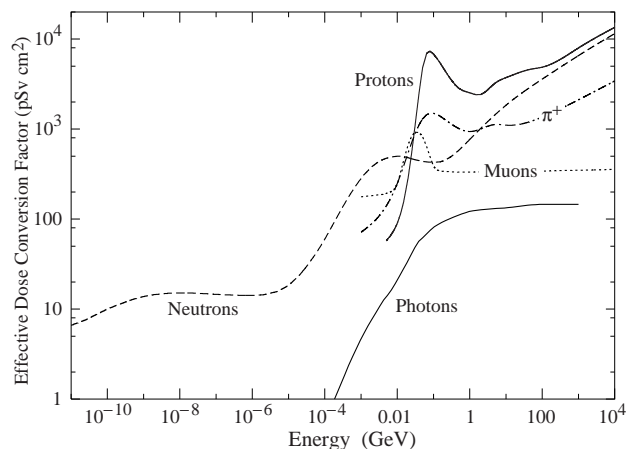
#### 36.1.4. Dose conversion coefficients :

Dose conversion coefficients allow direct calculation of protection or operational quantities from particle fluence and are functions of particle type, energy and irradiation configuration. The most common coefficients are those for effective dose and ambient dose equivalent. The former are based on simulations in which the dose to organs of anthropomorphic phantoms is calculated for approximate actual conditions of exposure, such as irradiation of the front of the body (anterior-posterior irradiation) or isotropic irradiation.

Conversion coefficients from fluence to effective dose are given for anterior-posterior irradiation and various particles in Fig. 36.1 [3]. For example, the effective dose from an anterior-posterior irradiation in a field of 1-MeV neutrons with a fluence of 1 neutron per  $\text{cm}^2$  is about 290 pSv. In Monte Carlo simulations such coefficients allow multiplication with fluence at scoring time such that effective dose to a human body at the considered location is directly obtained.

### 36.2. Radiation levels [4]

• **Natural background radiation**: On a worldwide average, the annual whole-body dose equivalent due to all sources of natural background radiation ranges from 1.0 to 13 mSv (0.1–1.3 rem) with an annual average of 2.4 mSv [5]. In certain areas values up to 50 mSv (5 rem) have been measured. A large fraction (typically more than 50%) originates from inhaled natural radioactivity, mostly radon and radon daughters. The latter can vary by more than one order of magnitude: it is 0.1–0.2 mSv in open areas, 2 mSv on average in a house and more than 20 mSv in poorly ventilated mines.



**Figure 36.1:** Fluence to effective dose conversion coefficients for anterior-posterior irradiation and various particles [3].

- **Cosmic ray background radiation:** At sea level, the whole-body dose equivalent due to cosmic ray background radiation is dominated by muons; at higher altitudes also nucleons contribute. Dose equivalent rates range from less than  $0.1 \mu\text{Sv/h}$  at sea level to a few  $\mu\text{Sv/h}$  at aircraft altitudes. Details on cosmic ray fluence levels are given in the Cosmic Rays section (Sec. 29 of this Review).
- **Fluence to deposit one Gy:** *Charged particles:* The fluence necessary to deposit a dose of one Gy (in units of  $\text{cm}^{-2}$ ) is about  $6.24 \times 10^9 / (dE/dx)$ , where  $dE/dx$  (in units of  $\text{MeV g}^{-1} \text{cm}^2$ ) is the mean energy loss rate that may be obtained from Figs. 33.2 and 33.4 in Sec. 33 of this Review, and from <http://pdg.lbl.gov/AtomicNuclearProperties>. For example, it is approximately  $3.5 \times 10^9 \text{ cm}^{-2}$  for minimum-ionizing singly-charged particles in carbon. *Photons:* This fluence is about  $6.24 \times 10^9 / (Ef/\ell)$  for photons of energy  $E$  (in MeV), an attenuation length  $\ell$  (in  $\text{g cm}^{-2}$ ), and a fraction  $f \lesssim 1$ , expressing the fraction of the photon energy deposited in a small volume of thickness  $\ll \ell$  but large enough to contain the secondary electrons. For example, it is approximately  $2 \times 10^{11} \text{ cm}^{-2}$  for 1 MeV photons on carbon ( $f \approx 1/2$ ).

### 36.3. Health effects of ionizing radiation

Radiation can cause two types of health effects, deterministic and stochastic:

- **Deterministic effects** are tissue reactions which cause injury to a population of cells if a given threshold of absorbed dose is exceeded. The severity of the reaction increases with dose. The quantity in use for tissue reactions is the absorbed dose,  $D$ . When particles other than photons and electrons (low-LET radiation) are involved, a Relative Biological Effectiveness (*RBE*)-weighted dose may be used. The *RBE* of a given radiation is the reciprocal of the ratio of the absorbed dose of that radiation to the absorbed dose of a reference radiation (usually X rays) required to produce the same degree of biological effect. It is a complex quantity that depends on many factors such as cell type, dose rate, fractionation, etc.
- **Stochastic effects** are malignant diseases and heritable effects for which the probability of an effect occurring, but not its severity, is a function of dose without threshold.
- **Lethal dose:** The whole-body dose from penetrating ionizing radiation resulting in 50% mortality in 30 days (assuming no medical treatment) is 2.5–4.5 Gy (250–450 rad)<sup>†</sup>, as measured internally on the body longitudinal center line. The surface dose varies due to variable body attenuation and may be a strong function of energy.
- **Cancer induction:** The cancer induction probability is about 5% per Sv on average for the entire population [2].
- **Recommended effective dose limits:** The International Commission on Radiological Protection (ICRP) recommends a limit for radiation workers of 20 mSv effective dose per year averaged over

<sup>†</sup> *RBE*-weighted when necessary

5 years, with the provision that the dose should not exceed 50 mSv in any single year [2]. The limit in the EU-countries and Switzerland is 20 mSv per year, in the U.S. it is 50 mSv per year (5 rem per year). Many physics laboratories in the U.S. and elsewhere set lower limits. The effective dose limit for general public is typically 1 mSv per year.

### 36.4. Prompt neutrons at accelerators

Neutrons dominate the particle environment outside thick shielding (*e.g.*,  $> 1 \text{ m}$  of concrete) for high energy ( $> \text{a few hundred MeV}$ ) electron and hadron accelerators. In addition, for accelerators with energies above about 10 GeV, muons contribute significantly at small angles with regard to the beam, even behind several meters of shielding. Another special case are synchrotron light sources where particular care has to be taken to shield the very intense low-energy photons extracted from the electron synchrotron into the experimental areas. Due to its importance at high energy accelerators this section focuses on prompt neutrons.

#### 36.4.1. Electron accelerators :

At electron accelerators, neutrons are generated via photonuclear reactions from bremsstrahlung photons. Neutron production takes place above a threshold value which varies from 10 to 19 MeV for light nuclei (with important exceptions, such as 2.23 MeV for deuterium and 1.67 MeV for beryllium) and from 4 to 6 MeV for heavy nuclei. It is commonly described by different mechanisms depending on the photon energy: the giant dipole resonance interactions (from threshold up to about 30 MeV, often the dominant process), the quasi-deuteron effect (between 30 MeV and a few hundred MeV), the delta resonance mechanism (between 200 MeV and a few GeV) and the vector meson dominance model at higher energies.

The giant dipole resonance reaction consists in a collective excitation of the nucleus, in which neutrons and protons oscillate in the direction of the photon electric field. The oscillation is damped by friction in a few cycles, with the photon energy being transferred to the nucleus in a process similar to evaporation. Nucleons emitted in the dipolar interaction have an anisotropic angular distribution, with a maximum at  $90^\circ$ , while those leaving the nucleus as a result of evaporation are emitted isotropically with a Maxwellian energy distribution described as [6]:

$$\frac{dN}{dE_n} = \frac{E_n}{T^2} e^{-E_n/T}, \quad (36.4)$$

where  $T$  is a nuclear ‘temperature’ (in units of MeV) characteristic of the particular target nucleus and its excitation energy. For heavy nuclei the ‘temperature’ generally lies in the range of  $T = 0.5\text{--}1.0$  MeV. Neutron yields from semi-infinite targets per kW of electron beam power are plotted in Fig. 36.2 as a function of the electron beam energy [6].

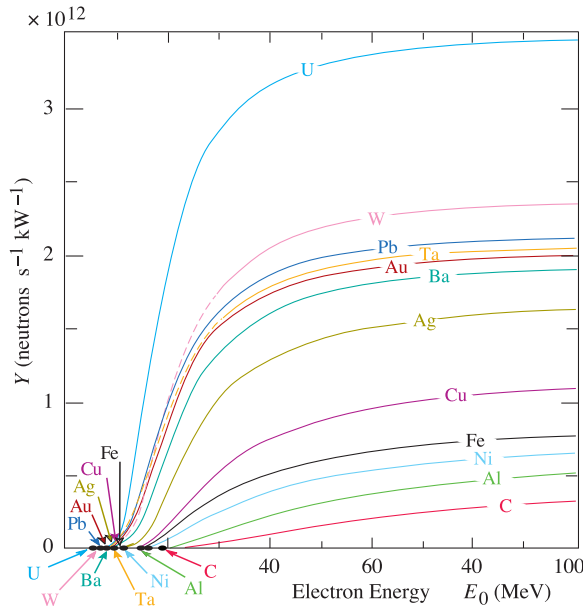
Typical neutron energy spectra outside of concrete (80 cm thick,  $2.35 \text{ g/cm}^3$ ) and iron (40 cm thick) shields are shown in Fig. 36.3. In order to compare these spectra to those caused by proton beams (see below) the spectra are scaled by a factor of 100, which roughly corresponds to the difference in the high energy hadronic cross sections for photons and hadrons (*e.g.*, the fine structure constant). The shape of these spectra are generally characterized by a low-energy peak at around 1 MeV (evaporation neutrons) and a high-energy shoulder at around 70–80 MeV. In case of concrete shielding, the spectrum also shows a pronounced peak at thermal neutron energies.

#### 36.4.2. Proton accelerators :

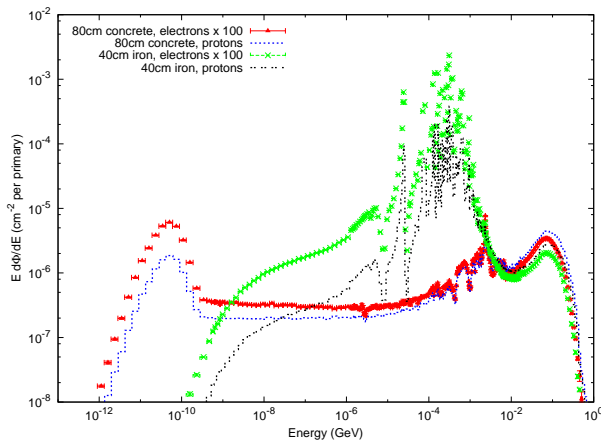
At proton accelerators, neutron yields emitted per incident proton by different target materials are roughly independent of proton energy between 20 MeV and 1 GeV, and are given by the ratio C : Al : Cu-Fe : Sn : Ta-Pb = 0.3 : 0.6 : 1.0 : 1.5 : 1.7 [9]. Above about 1 GeV, the neutron yield is proportional to  $E^m$ , where  $0.80 \leq m \leq 0.85$  [10].

Typical neutron energy spectra outside of concrete and iron shielding are shown in Fig. 36.3. Here, the radiation fields are caused by a 25 GeV proton beam interacting with a thick copper target. The comparison of these spectra with those for an electron beam of the same energy reflects the difference in the hadronic cross sections





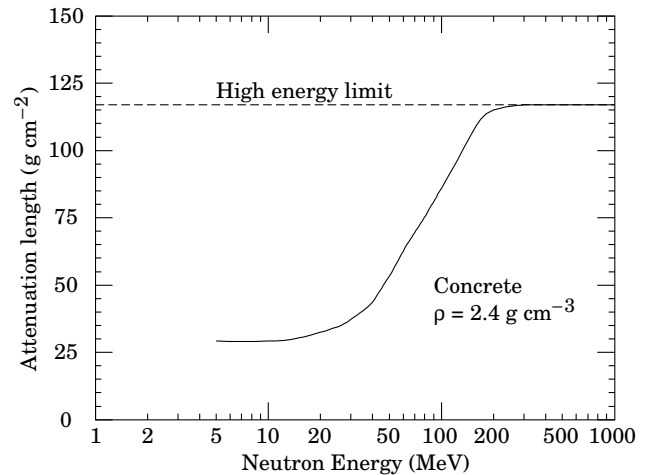
**Figure 36.2:** Neutron yields from semi-infinite targets per kW of electron beam power, as a function of the electron beam energy, disregarding target self-shielding [6].



**Figure 36.3:** Neutron energy spectra calculated with the FLUKA code [7,8] from 25 GeV proton and electron beams on a thick copper target. Spectra are evaluated at  $90^\circ$  to the beam direction behind 80 cm of concrete or 40 cm of iron. All spectra are normalized per beam particle. In addition, spectra for electron beam are multiplied by a factor of 100.

between photons and hadrons above a few 100 MeV. Differences are increasing towards lower energies because of different interaction mechanisms. Furthermore, the slight shift in energy above about 100 MeV follows from the fact that the energies of the interacting photons are lower than 25 GeV. Apart from this the shapes of the two spectra are similar.

The neutron-attenuation length is shown in Fig. 36.4 for concrete and mono-energetic broad-beam conditions. As can be seen in the figure it reaches a value of about  $117 \text{ g/cm}^2$  above 200 MeV. As the cascade through thick shielding is carried by high-energy particles this value is equal to the equilibrium attenuation length for particles emitted at  $90^\circ$  in concrete.

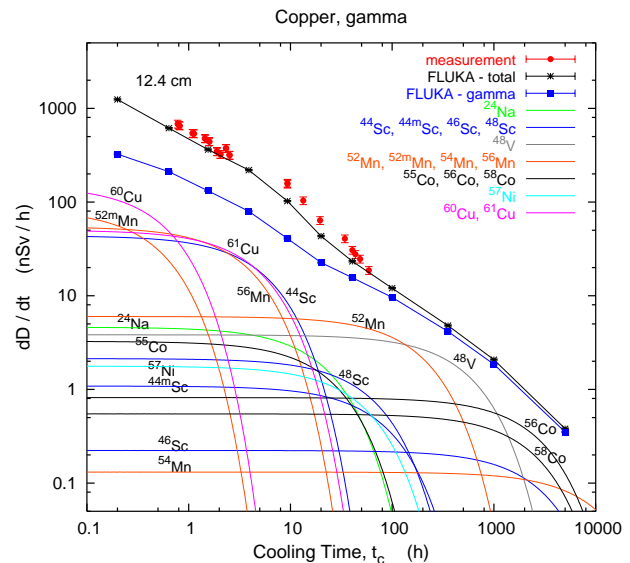


**Figure 36.4:** The variation of the attenuation length for mono-energetic neutrons in concrete as a function of neutron energy [9].

### 36.5. Photon sources

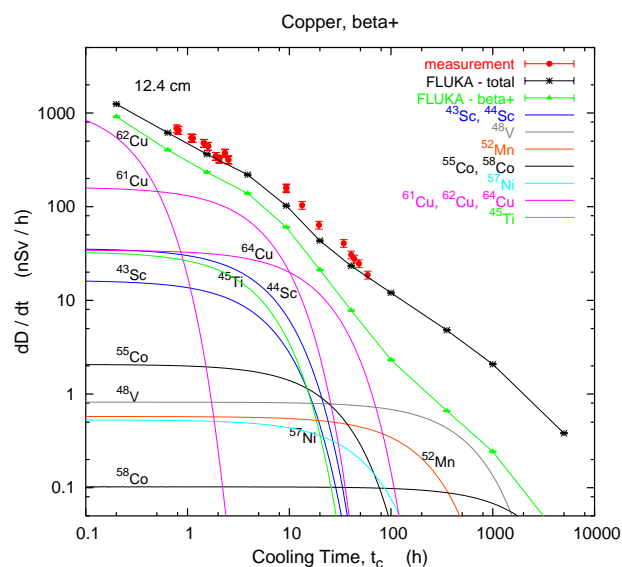
The dose equivalent rate in tissue (in mSv/h) from a gamma point source emitting one photon of energy  $E$  (in MeV) per second at a distance of 1 m is  $4.6 \times 10^{-9} \mu_{en}/\rho E$ , where  $\mu_{en}/\rho$  is the mass energy absorption coefficient. The latter has a value of  $0.029 \pm 0.004 \text{ cm}^2/\text{g}$  for photons in tissue over an energy range between 60 keV and 2 MeV (see Ref. 11 for tabulated values).

Similarly, the dose equivalent rate in tissue (in mSv/h) at the surface of a semi-infinite slab of uniformly activated material containing 1 Bq/g of a gamma emitter of energy  $E$  (in MeV) is  $2.9 \times 10^{-4} R_\mu E$ , where  $R_\mu$  is the ratio of the mass energy absorption coefficients of the photons in tissue and in the material.



**Figure 36.5:** Contribution of individual gamma-emitting nuclides to the total dose rate at 12.4 cm distance to an activated copper sample [12].





**Figure 36.6:** Contribution of individual positron-emitting nuclides to the total dose rate at 12.4 cm distance to an activated copper sample [12].

### 36.6. Accelerator-induced radioactivity

Typical medium- and long-lived activation products in metallic components of accelerators are  $^{22}\text{Na}$ ,  $^{46}\text{Sc}$ ,  $^{48}\text{V}$ ,  $^{51}\text{Cr}$ ,  $^{54}\text{Mn}$ ,  $^{55}\text{Fe}$ ,  $^{59}\text{Fe}$ ,  $^{56}\text{Co}$ ,  $^{57}\text{Co}$ ,  $^{58}\text{Co}$ ,  $^{60}\text{Co}$ ,  $^{63}\text{Ni}$  and  $^{65}\text{Zn}$ . Gamma-emitting nuclides dominate doses by external irradiation at longer decay times (more than one day) while at short decay times  $\beta^+$  emitters are also important (through photons produced by  $\beta^+$  annihilation). Due to their short range,  $\beta^-$  emitters are relevant, for example, only for dose to the skin and eyes or for doses due to inhalation or ingestion. Fig. 36.5 and Fig. 36.6 show the contributions of gamma and  $\beta^+$  emitters to the total dose rate at 12.4 cm distance to a copper sample [12]. The sample was activated by the stray radiation field created by a 120 GeV mixed hadron beam dumped in a copper target during about 8 hours at intensities between  $10^7 - 10^8$  hadrons per second. Typically, dose rates at a certain decay time are mainly determined by radionuclides having a half-life of the order of the decay time. Extended irradiation periods might be an exception to this general rule as in this case the activity of long-lived nuclides can build up sufficiently so that it dominates that one of short-lived even at short cooling times.

Activation in concrete is dominated by  $^{24}\text{Na}$  (short decay times) and  $^{22}\text{Na}$  (long decay times). Both nuclides can be produced either by low-energy neutron reactions on the sodium-component in the concrete or by spallation reactions on silicon, calcium and other constituents such as aluminum. At long decay times nuclides of radiological interest in activated concrete can also be  $^{60}\text{Co}$ ,  $^{152}\text{Eu}$ ,  $^{154}\text{Eu}$  and  $^{134}\text{Cs}$ , all of which produced by  $(n,\gamma)$ -reactions with traces of natural cobalt, europium and cesium. Thus, such trace elements might be important even if their content in concrete is only a few parts per million or less by weight.

The explicit simulation of radionuclide production with general-purpose Monte Carlo codes has become the most commonly applied method to calculate induced radioactivity and its radiological consequences. Nevertheless, other more approximative approaches, such as “ $\omega$ -factors” [9], can still be useful for fast order-of-magnitude estimates. These  $\omega$ -factors give the dose rate per unit star density (inelastic reactions above a certain energy threshold, *e.g.* 50 MeV) on contact to an extended, uniformly activated object after a 30-day irradiation and 1-day decay. For steel or iron,  $\omega \approx 3 \times 10^{-12}$  (Sv  $\text{cm}^3/\text{star}$ ). This does not include possible contributions from thermal-neutron activation.

### 36.7. Radiation protection instrumentation

The capacity to distinguish and measure the high-LET (mostly neutrons) and the low-LET components (photons, electrons, muons) of the radiation field at workplaces is of primary importance to evaluate the exposure of personnel. At proton machines the prompt dose equivalent outside a shield is mainly due to neutrons, with some contribution from photons and, to a minor extent, charged particles. At high-energy electron accelerators the dominant stray radiation during operation consists of high-energy neutrons, because the shielding is normally thick enough to absorb most of the bremsstrahlung photons. Most of the personnel exposure at accelerator facilities is often received during maintenance interventions, and is due to gamma/beta radiation coming from residual radioactivity in accelerator components.

Radiation detectors used both for radiation surveys and area monitoring are normally calibrated in ambient dose equivalent  $H^*(10)$ .

#### 36.7.1. Neutron detectors :

- **Rem counters:** A rem counter is a portable detector consisting of a thermal neutron counter embedded in a polyethylene moderator, with a response function that approximately follows the curve of the conversion coefficients from neutron fluence to  $H^*(10)$  over a wide energy range. Conventional rem counters provide a response to neutrons up to approximately 10-15 MeV, extended-range units are heavier as they include a high-Z converter but correctly measure  $H^*(10)$  up to several hundred MeV.

- **Bonner Sphere Spectrometer (BSS):** A BSS is made up of a thermal neutron detector at the centre of moderating spheres of different diameters made of polyethylene (PE) or a combination of PE and a high-Z material. Each sphere has a different response function versus neutron energy, and the neutron energy, at which the sensitivity peaks, increases with sphere diameter. The energy resolution of the system is rather low but satisfactory for radiation protection purposes. The neutron spectrum is obtained by unfolding the experimental counts of the BSS with its response matrix by a computer code that is often based on an iterative algorithm. BSS exist in active (using  $^3\text{He}$  or  $\text{BF}_3$  proportional counters or  $^6\text{LiI}$  scintillators) and passive versions (using CR-39 track detectors or LiF), for use *e.g.* in strongly pulsed fields. With  $^3\text{He}$  counters the discrimination with respect to gamma rays and noise is excellent.

- **Bubble detectors:** A bubble detector is a dosimeter based on a super-heated emulsion (super-heated droplets suspended in a gel) contained in a vial and acting as a continuously sensitive, miniature bubble chamber. The total number of bubbles evolved from the radiation-induced nucleation of drops gives an integrated measure of the total neutron exposure. Various techniques exist to record and count the bubbles, *e.g.*, visual inspection, automated reading with video cameras or acoustic counting. Bubble detectors are insensitive to low-LET radiation. Super-heated emulsions are used as personal, area and environmental dosimeters, as well as neutron spectrometers.

- **Track etched detectors:** Track etched detectors (TEDs) are based on the preferential dissolution of suitable, mostly insulator, materials along the damage trails of charged particles of sufficiently high-energy deposition density. The detectors are effectively not sensitive to radiation which deposits the energy through the interactions of particles with low LET. These dosimeters are generally able to determine neutron ambient dose equivalent down to around 100  $\mu\text{Sv}$ . They are used both as personal dosimeters and for area monitoring, *e.g.*, in BSS.

#### 36.7.2. Photon detectors :

- **GM counters:** Geiger Müller (GM) counters are low cost devices and simple to operate. They work in pulse mode and since they only count radiation-induced events, any spectrometric information is lost. In general they are calibrated in terms of air kerma, for instance in a  $^{60}\text{Co}$  field. The response of GM counters to photons is constant within 15% for energies up to 2 MeV and shows considerable energy dependence above.

- **Ionization chambers:** Ionization chambers are gas-filled detectors used both as hand-held instruments (*e.g.*, for radiation surveys)

and environmental monitors. They are normally operated in current mode although pulse-mode operation is also possible. They possess a relatively flat response to a wide range of X- and gamma ray energies (typically from 10 keV to several MeV), can measure radiation over a wide intensity range and are capable of discriminating between the beta and gamma components of a radiation field (by use of, *e.g.*, a beta window). Pressurized ion chambers (filled, *e.g.*, with Ar or H gas to several tens of bars) are used for environmental monitoring applications. They have good sensitivity to neutrons and charged hadrons in addition to low LET radiation (gammas and muons), with the response function to the former being strongly non-linear with energy.

• **Scintillators:** Scintillation-based detectors are used in radiation protection as hand-held probes and in fixed installations, *e.g.*, portal monitors. A scintillation detector or counter is obtained coupling a scintillator to an electronic light sensor such as a photomultiplier tube (PMT), a photodiode or a silicon photomultiplier (SiPM). There is a wide range of scintillating materials, inorganic (such as CsI and BGO), organic or plastic; they find application in both photon dosimetry and spectrometry.

### 36.7.3. Personal dosimeters :

Personal dosimeters, calibrated in  $H_p(10)$ , are worn by persons exposed to ionizing radiation for professional reasons to record the dose received. They are typically passive detectors, either film, track etched detectors,  $^6\text{Li}/^7\text{Li}$ -based dosimeters (*e.g.* LiF), optically stimulated luminescence (OSL) or radiophotoluminescence detectors (RPL) but semi-active dosimeters using miniaturized ion-chambers also exist.

Electronic personal dosimeters are small active units for on-line monitoring of individual exposure, designed to be worn on the body. They can give an alarm on both the integral dose received or dose rate once a pre-set threshold is exceeded.

## 36.8. Monte Carlo codes for radiation protection studies

The use of general-purpose particle interaction and transport Monte Carlo codes is often the most accurate and efficient choice for assessing radiation protection quantities at accelerators. Due to the vast spread of such codes to all areas of particle physics and the associated extensive benchmarking with experimental data, the modeling has reached an unprecedented accuracy. Furthermore, most codes allow the user to simulate all aspects of a high energy particle cascade in one and the same run: from the first interaction of a TeV nucleus over the transport and re-interactions (hadronic and electromagnetic) of the produced secondaries, to detailed nuclear fragmentation, the calculation of radioactive decays and even of the electromagnetic shower caused by the radiation from such decays. A brief account of the codes most widely used for radiation protection studies at high energy accelerators is given in the following.

• **FLUKA [7,8]:** FLUKA is a general-purpose particle interaction and transport code. It comprises all features needed for radiation protection, such as detailed hadronic and nuclear interaction models up to 10 PeV, full coupling between hadronic and electromagnetic processes and numerous variance reduction options. The latter include weight windows, region importance biasing, and leading particle, interaction, and decay length biasing (among others). The capabilities of FLUKA are unique for studies of induced radioactivity, especially with regard to nuclide production, decay, and transport of residual radiation. In particular, particle cascades by prompt and residual radiation are simulated in parallel based on the microscopic models for nuclide production and a solution of the Bateman equations for activity build-up and decay.

• **GEANT4 [13,14]:** GEANT4 is an object-oriented toolkit consisting of a kernel that provides the framework for particle transport, including tracking, geometry description, material specifications, management of events and interfaces to external graphics systems. The kernel also provides interfaces to physics processes. It allows the user to freely select the physics models that best serve the particular application needs. Implementations of interaction models exist over an extended range of energies, from optical photons and thermal

neutrons to high-energy interactions required for the simulation of accelerator and cosmic ray experiments. To facilitate the use of variance reduction techniques, general-purpose biasing methods such as importance biasing, weight windows, and a weight cut-off method have been introduced directly into the toolkit. Other variance reduction methods, such as leading particle biasing for hadronic processes, come with the respective physics packages.

• **MARS15 [15,16]:** The MARS15 code system is a set of Monte Carlo programs for the simulation of hadronic and electromagnetic cascades. It covers a wide energy range: 1 keV to 100 TeV for muons, charged hadrons, heavy ions and electromagnetic showers; and 0.00215 eV to 100 TeV for neutrons. Hadronic interactions above 5 GeV can be simulated with either an inclusive or an exclusive event generator. MARS15 is coupled to the MCNP4C code that handles all interactions of neutrons with energies below 14 MeV. Different variance reduction techniques, such as inclusive particle production, weight windows, particle splitting, and Russian roulette, are available in MARS15. A tagging module allows one to tag the origin of a given signal for source term or sensitivity analyses. Further features of MARS15 include a MAD-MARS Beam-Line Builder for a convenient creation of accelerator models.

• **MCNPX [17,18]:** MCNPX originates from the Monte Carlo N-Particle transport (MCNP) family of neutron interaction and transport codes and, therefore, features one of the most comprehensive and detailed descriptions of the related physical processes. Later it was extended to other particle types, including ions and electromagnetic particles. The neutron interaction and transport modules use standard evaluated data libraries mixed with physics models where such libraries are not available. The transport is continuous in energy. MCNPX contains one of the most powerful implementations of variance reduction techniques. Spherical mesh weight windows can be created by a generator in order to focus the simulation time on certain spatial regions of interest. In addition, a more generalized phase space biasing is also possible through energy- and time-dependent weight windows. Other biasing options include pulse-height tallies with variance reduction and criticality source convergence acceleration.

• **PHITS [19,20]:** The Particle and Heavy-Ion Transport code System PHITS was among the first general-purpose codes to simulate the transport and interactions of heavy ions in a wide energy range, from 10 MeV/nucleon to 100 GeV/nucleon. It is based on the high-energy hadron transport code NMTC/JAM that was extended to heavy ions. The transport of low-energy neutrons employs cross sections from evaluated nuclear data libraries such as ENDF and JENDL below 20 MeV and LA150 up to 150 MeV. Electromagnetic interactions are simulated based on the ITS code in the energy range between 1 keV and 1 GeV. Several variance reduction techniques, including weight windows and region importance biasing, are available in PHITS.

### References:

1. International Commission on Radiation Units and Measurements, *Fundamental Quantities and Units for Ionizing Radiation*, ICRU Report 60 (1998).
2. ICRP Publication 103, *The 2007 Recommendations of the International Commission on Radiological Protection*, Annals of the ICRP, Elsevier (2007).
3. M. Pelliccioni, *Radiation Protection Dosimetry* **88**, 279 (2000).
4. E. Pochin, *Nuclear Radiation: Risks and Benefits*, Clarendon Press, Oxford, 1983.
5. United Nations, *Report of the United Nations Scientific Committee on the Effect of Atomic Radiation*, General Assembly, Official Records A/63/46 (2008).
6. W.P. Swanson, *Radiological Safety Aspects of the Operation of Electron Linear Accelerators*, IAEA Technical Reports Series No. 188 (1979).
7. A. Ferrari, *et al.*, FLUKA, A Multi-particle Transport Code (Program Version 2005), CERN-2005-010 (2005).
8. G. Battistoni, *et al.*, The FLUKA code: Description and benchmarking, *Proceedings of the Hadronic Shower Simulation Workshop 2006*, Fermilab 6–8 September 2006, M. Albrow, R. Raja, eds., *AIP Conference Proceeding 896*, 31–49, (2007).

9. R.H. Thomas and G.R. Stevenson, *Radiological Safety Aspects of the Operation of Proton Accelerators*, IAEA Technical Report Series No. 283 (1988).
10. T.A. Gabriel, *et al.*, Nucl. Instrum. Methods **A338**, 336 (1994).
11. <http://physics.nist.gov/PhysRefData/XrayMassCoef/cover.html>.
12. S. Roesler, *et al.*, "Simulation of Remanent Dose Rates and Benchmark Measurements at the CERN-EU High Energy Reference Field Facility," in *Proceedings of the Sixth International Meeting on Nuclear Applications of Accelerator Technology*, San Diego, CA, 1-5 June 2003, 655-662 (2003).
13. S. Agostinelli, *et al.*, Nucl. Instrum. Methods **A506**, 250 (2003).
14. J. Allison, *et al.*, IEEE Transactions on Nuclear Science **53**, 270 (2006).
15. N.V. Mokhov, S.I. Striganov, MARS15 Overview *Proceedings of the Hadronic Shower Simulation Workshop 2006*, Fermilab 6-8 September 2006, M. Albrow, R. Raja, eds., *AIP Conference Proceeding 896*, 50-60, (2007).
16. N.V. Mokhov, MARS Code System, Version 15 (2009), [www-ap.fnal.gov/MARS](http://www-ap.fnal.gov/MARS).
17. D.B. Pelowitz, ed., Los Alamos National Laboratory report, LA-CP-05-0369 (2005).
18. G. McKinney, *et al.*, *Proceedings of the International Workshop on Fast Neutron Detectors* University of Cape Town, South Africa (2006).
19. H. Iwase, K. Niita, and T. Nakamura, Journal of Nuclear Science and Technology **39**, 1142 (2002).
20. K. Niita, *et al.*, Radiation Measurements **41**, 1080 (2006).

37. COMMONLY USED RADIOACTIVE SOURCES

Table 37.1. Revised November 1993 by E. Browne (LBNL).

Nuclide	Half-life	Type of decay	Particle		Photon	
			Energy (MeV)	Emission prob.	Energy (MeV)	Emission prob.
<sup>22</sup> <sub>11</sub> Na	2.603 y	$\beta^+$ , EC	0.545	90%	0.511 Annih. 1.275 100%	
<sup>54</sup> <sub>25</sub> Mn	0.855 y	EC			0.835 100% Cr K x rays 26%	
<sup>55</sup> <sub>26</sub> Fe	2.73 y	EC			Mn K x rays: 0.00590 24.4% 0.00649 2.86%	
<sup>57</sup> <sub>27</sub> Co	0.744 y	EC			0.014 9% 0.122 86% 0.136 11% Fe K x rays 58%	
<sup>60</sup> <sub>27</sub> Co	5.271 y	$\beta^-$	0.316	100%	1.173 100% 1.333 100%	
<sup>68</sup> <sub>32</sub> Ge	0.742 y	EC			Ga K x rays 44%	
-----						
$\rightarrow$ <sup>68</sup> <sub>31</sub> Ga		$\beta^+$ , EC	1.899	90%	0.511 Annih. 1.077 3%	
<sup>90</sup> <sub>38</sub> Sr	28.5 y	$\beta^-$	0.546	100%		
-----						
$\rightarrow$ <sup>90</sup> <sub>39</sub> Y		$\beta^-$	2.283	100%		
<sup>106</sup> <sub>44</sub> Ru	1.020 y	$\beta^-$	0.039	100%		
-----						
$\rightarrow$ <sup>106</sup> <sub>45</sub> Rh		$\beta^-$	3.541	79%	0.512 21% 0.622 10%	
<sup>109</sup> <sub>48</sub> Cd	1.267 y	EC	0.063 $e^-$ 0.084 $e^-$ 0.087 $e^-$	41% 45% 9%	0.088 3.6% Ag K x rays 100%	
<sup>113</sup> <sub>50</sub> Sn	0.315 y	EC	0.364 $e^-$ 0.388 $e^-$	29% 6%	0.392 65% In K x rays 97%	
<sup>137</sup> <sub>55</sub> Cs	30.2 y	$\beta^-$	0.514 1.176	94% 6%	0.662 85%	
-----						
<sup>133</sup> <sub>56</sub> Ba	10.54 y	EC	0.045 $e^-$ 0.075 $e^-$	50% 6%	0.081 34% 0.356 62% Cs K x rays 121%	
<sup>207</sup> <sub>83</sub> Bi	31.8 y	EC	0.481 $e^-$ 0.975 $e^-$ 1.047 $e^-$	2% 7% 2%	0.569 98% 1.063 75% 1.770 7% Pb K x rays 78%	
<sup>228</sup> <sub>90</sub> Th	1.912 y	6 $\alpha$ : 3 $\beta^-$ :	5.341 to 8.785 0.334 to 2.246		0.239 44% 0.583 31% 2.614 36%	
( $\rightarrow$ <sup>224</sup> <sub>88</sub> Ra $\rightarrow$ <sup>220</sup> <sub>86</sub> Rn $\rightarrow$ <sup>216</sup> <sub>84</sub> Po $\rightarrow$ <sup>212</sup> <sub>82</sub> Pb $\rightarrow$ <sup>212</sup> <sub>83</sub> Bi $\rightarrow$ <sup>212</sup> <sub>84</sub> Po )						
<sup>241</sup> <sub>95</sub> Am	432.7 y	$\alpha$	5.443 5.486	13% 85%	0.060 36% Np L x rays 38%	
<sup>241</sup> <sub>95</sub> Am/Be	432.2 y	6 $\times$ 10 <sup>-5</sup> neutrons (4-8 MeV) and 4 $\times$ 10 <sup>-5</sup> $\gamma$ 's (4.43 MeV) per Am decay				
<sup>244</sup> <sub>96</sub> Cm	18.11 y	$\alpha$	5.763 5.805	24% 76%	Pu L x rays $\sim$ 9%	
<sup>252</sup> <sub>98</sub> Cf	2.645 y	$\alpha$ (97%) Fission (3.1%)	6.076 6.118	15% 82%		
$\approx$ 20 $\gamma$ 's/fission; 80% < 1 MeV $\approx$ 4 neutrons/fission; $\langle E_n \rangle$ = 2.14 MeV						

“Emission probability” is the probability per decay of a given emission; because of cascades these may total more than 100%. Only principal emissions are listed. EC means electron capture, and  $e^-$  means monoenergetic internal conversion (Auger) electron. The intensity of 0.511 MeV  $e^+e^-$  annihilation photons depends upon the number of stopped positrons. Endpoint  $\beta^\pm$  energies are listed. In some cases when energies are closely spaced, the  $\gamma$ -ray values are approximate weighted averages. Radiation from short-lived daughter isotopes is included where relevant.

Half-lives, energies, and intensities are from E. Browne and R.B. Firestone, *Table of Radioactive Isotopes* (John Wiley & Sons, New York, 1986), recent *Nuclear Data Sheets*, and *X-ray and Gamma-ray Standards for Detector Calibration*, IAEA-TECDOC-619 (1991).

Neutron data are from *Neutron Sources for Basic Physics and Applications* (Pergamon Press, 1983).

## 38. PROBABILITY

Revised September 2015 by G. Cowan (RHUL).

### 38.1. General [1–8]

An abstract definition of probability can be given by considering a set  $S$ , called the sample space, and possible subsets  $A, B, \dots$ , the interpretation of which is left open. The probability  $P$  is a real-valued function defined by the following axioms due to Kolmogorov [9]:

1. For every subset  $A$  in  $S$ ,  $P(A) \geq 0$ ;
2. For disjoint subsets (*i.e.*,  $A \cap B = \emptyset$ ),  $P(A \cup B) = P(A) + P(B)$ ;
3.  $P(S) = 1$ .

In addition, one defines the conditional probability  $P(A|B)$  (read as  $P$  of  $A$  given  $B$ ) as

$$P(A|B) = \frac{P(A \cap B)}{P(B)}. \quad (38.1)$$

From this definition and using the fact that  $A \cap B$  and  $B \cap A$  are the same, one obtains *Bayes' theorem*,

$$P(A|B) = \frac{P(B|A)P(A)}{P(B)}. \quad (38.2)$$

From the three axioms of probability and the definition of conditional probability, one obtains the *law of total probability*,

$$P(B) = \sum_i P(B|A_i)P(A_i), \quad (38.3)$$

for any subset  $B$  and for disjoint  $A_i$  with  $\cup_i A_i = S$ . This can be combined with Bayes' theorem (Eq. (38.2)) to give

$$P(A|B) = \frac{P(B|A)P(A)}{\sum_i P(B|A_i)P(A_i)}, \quad (38.4)$$

where the subset  $A$  could, for example, be one of the  $A_i$ .

The most commonly used interpretation of the elements of the sample space are outcomes of a repeatable experiment. The probability  $P(A)$  is assigned a value equal to the limiting frequency of occurrence of  $A$ . This interpretation forms the basis of *frequentist statistics*.

The elements of the sample space might also be interpreted as *hypotheses*, *i.e.*, statements that are either true or false, such as ‘The mass of the  $W$  boson lies between 80.3 and 80.5 GeV.’ Upon repetition of a measurement, however, such statements are either always true or always false, *i.e.*, the corresponding probabilities in the frequentist interpretation are either 0 or 1. Using *subjective probability*, however,  $P(A)$  is interpreted as the degree of belief that the hypothesis  $A$  is true. Subjective probability is used in *Bayesian* (as opposed to frequentist) statistics. Bayes' theorem can be written

$$P(\text{theory}|\text{data}) \propto P(\text{data}|\text{theory})P(\text{theory}), \quad (38.5)$$

where ‘theory’ represents some hypothesis and ‘data’ is the outcome of the experiment. Here  $P(\text{theory})$  is the *prior* probability for the theory, which reflects the experimenter's degree of belief before carrying out the measurement, and  $P(\text{data}|\text{theory})$  is the probability to have gotten the data actually obtained, given the theory, which is also called the *likelihood*.

Bayesian statistics provides no fundamental rule for obtaining the prior probability, which may depend on previous measurements, theoretical prejudices, *etc.* Once this has been specified, however, Eq. (38.5) tells how the probability for the theory must be modified in the light of the new data to give the *posterior* probability,  $P(\text{theory}|\text{data})$ . As Eq. (38.5) is stated as a proportionality, the probability must be normalized by summing (or integrating) over all possible hypotheses.

### 38.2. Random variables

A *random variable* is a numerical characteristic assigned to an element of the sample space. In the frequency interpretation of probability, it corresponds to an outcome of a repeatable experiment. Let  $x$  be a possible outcome of an observation. If  $x$  can take on any value from a continuous range, we write  $f(x;\theta)dx$  as the probability that the measurement's outcome lies between  $x$  and  $x+dx$ . The function  $f(x;\theta)$  is called the *probability density function* (p.d.f.), which may depend on one or more parameters  $\theta$ . If  $x$  can take on only discrete values (*e.g.*, the non-negative integers), then we use  $f(x;\theta)$  to denote the probability to find the value  $x$ . In the following the term p.d.f. is often taken to cover both the continuous and discrete cases, although technically the term density should only be used in the continuous case.

The p.d.f. is always normalized to unity. Both  $x$  and  $\theta$  may have multiple components and are then often written as vectors. If  $\theta$  is unknown, we may wish to estimate its value from a given set of measurements of  $x$ ; this is a central topic of *statistics* (see Sec. 39).

The *cumulative distribution function*  $F(a)$  is the probability that  $x \leq a$ :

$$F(a) = \int_{-\infty}^a f(x) dx. \quad (38.6)$$

Here and below, if  $x$  is discrete-valued, the integral is replaced by a sum. The endpoint  $a$  is expressly included in the integral or sum. Then  $0 \leq F(x) \leq 1$ ,  $F(x)$  is nondecreasing, and  $P(a < x \leq b) = F(b) - F(a)$ . If  $x$  is discrete,  $F(x)$  is flat except at allowed values of  $x$ , where it has discontinuous jumps equal to  $f(x)$ .

Any function of random variables is itself a random variable, with (in general) a different p.d.f. The *expectation value* of any function  $u(x)$  is

$$E[u(x)] = \int_{-\infty}^{\infty} u(x) f(x) dx, \quad (38.7)$$

assuming the integral is finite. The expectation value is linear, *i.e.*, for any two functions  $u$  and  $v$  of  $x$  and constants  $c_1$  and  $c_2$ ,  $E[c_1 u + c_2 v] = c_1 E[u] + c_2 E[v]$ .

The  $n^{\text{th}}$  moment of a random variable  $x$  is

$$\alpha_n \equiv E[x^n] = \int_{-\infty}^{\infty} x^n f(x) dx, \quad (38.8a)$$

and the  $n^{\text{th}}$  central moment of  $x$  (or moment about the mean,  $\alpha_1$ ) is

$$m_n \equiv E[(x - \alpha_1)^n] = \int_{-\infty}^{\infty} (x - \alpha_1)^n f(x) dx. \quad (38.8b)$$

The most commonly used moments are the mean  $\mu$  and variance  $\sigma^2$ :

$$\mu \equiv \alpha_1, \quad (38.9a)$$

$$\sigma^2 \equiv V[x] \equiv m_2 = \alpha_2 - \mu^2. \quad (38.9b)$$

The mean is the location of the ‘center of mass’ of the p.d.f., and the variance is a measure of the square of its width. Note that  $V[cx+k] = c^2 V[x]$ . It is often convenient to use the *standard deviation* of  $x$ ,  $\sigma$ , defined as the square root of the variance.

Any odd moment about the mean is a measure of the skewness of the p.d.f. The simplest of these is the dimensionless coefficient of skewness  $\gamma_1 = m_3/\sigma^3$ .

The fourth central moment  $m_4$  provides a convenient measure of the tails of a distribution. For the Gaussian distribution (see Sec. 38.4), one has  $m_4 = 3\sigma^4$ . The *kurtosis* is defined as  $\gamma_2 = m_4/\sigma^4 - 3$ , *i.e.*, it is zero for a Gaussian, positive for a *leptokurtic* distribution with longer tails, and negative for a *platykurtic* distribution with tails that die off more quickly than those of a Gaussian.

The *quantile*  $x_\alpha$  is the value of the random variable  $x$  at which the cumulative distribution is equal to  $\alpha$ . That is, the quantile is the inverse of the cumulative distribution function, *i.e.*,  $x_\alpha = F^{-1}(\alpha)$ . An important special case is the *median*,  $x_{\text{med}}$ , defined by  $F(x_{\text{med}}) = 1/2$ , *i.e.*, half the probability lies above and half lies below  $x_{\text{med}}$ .

(More rigorously,  $x_{\text{med}}$  is a median if  $P(x \geq x_{\text{med}}) \geq 1/2$  and  $P(x \leq x_{\text{med}}) \geq 1/2$ . If only one value exists, it is called ‘the median.’)

Under a monotonic change of variable  $x \rightarrow y(x)$ , the quantiles of a distribution (and hence also the median) obey  $y_\alpha = y(x_\alpha)$ . In general the expectation value and *mode* (most probable value) of a distribution do not, however, transform in this way.

Let  $x$  and  $y$  be two random variables with a *joint* p.d.f.  $f(x, y)$ . The *marginal* p.d.f. of  $x$  (the distribution of  $x$  with  $y$  unobserved) is

$$f_1(x) = \int_{-\infty}^{\infty} f(x, y) dy, \quad (38.10)$$

and similarly for the marginal p.d.f.  $f_2(y)$ . The *conditional* p.d.f. of  $y$  given fixed  $x$  (with  $f_1(x) \neq 0$ ) is defined by  $f_3(y|x) = f(x, y)/f_1(x)$ , and similarly  $f_4(x|y) = f(x, y)/f_2(y)$ . From these, we immediately obtain Bayes’ theorem (see Eqs. (38.2) and (38.4)),

$$f_4(x|y) = \frac{f_3(y|x)f_1(x)}{f_2(y)} = \frac{f_3(y|x)f_1(x)}{\int f_3(y|x')f_1(x') dx'}. \quad (38.11)$$

The mean of  $x$  is

$$\mu_x = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} x f(x, y) dx dy = \int_{-\infty}^{\infty} x f_1(x) dx, \quad (38.12)$$

and similarly for  $y$ . The *covariance* of  $x$  and  $y$  is

$$\text{cov}[x, y] = E[(x - \mu_x)(y - \mu_y)] = E[xy] - \mu_x \mu_y. \quad (38.13)$$

A dimensionless measure of the covariance of  $x$  and  $y$  is given by the *correlation coefficient*,

$$\rho_{xy} = \text{cov}[x, y] / \sigma_x \sigma_y, \quad (38.14)$$

where  $\sigma_x$  and  $\sigma_y$  are the standard deviations of  $x$  and  $y$ . It can be shown that  $-1 \leq \rho_{xy} \leq 1$ .

Two random variables  $x$  and  $y$  are *independent* if and only if

$$f(x, y) = f_1(x)f_2(y). \quad (38.15)$$

If  $x$  and  $y$  are independent, then  $\rho_{xy} = 0$ ; the converse is not necessarily true. If  $x$  and  $y$  are independent,  $E[u(x)v(y)] = E[u(x)]E[v(y)]$ , and  $V[x + y] = V[x] + V[y]$ ; otherwise,  $V[x + y] = V[x] + V[y] + 2\text{cov}[x, y]$ , and  $E[uv]$  does not necessarily factorize.

Consider a set of  $n$  continuous random variables  $\mathbf{x} = (x_1, \dots, x_n)$  with joint p.d.f.  $f(\mathbf{x})$ , and a set of  $n$  new variables  $\mathbf{y} = (y_1, \dots, y_n)$ , related to  $\mathbf{x}$  by means of a function  $\mathbf{y}(\mathbf{x})$  that is one-to-one, *i.e.*, the inverse  $\mathbf{x}(\mathbf{y})$  exists. The joint p.d.f. for  $\mathbf{y}$  is given by

$$g(\mathbf{y}) = f(\mathbf{x}(\mathbf{y}))|J|, \quad (38.16)$$

where  $|J|$  is the absolute value of the determinant of the square matrix  $J_{ij} = \partial x_i / \partial y_j$  (the Jacobian determinant). If the transformation from  $\mathbf{x}$  to  $\mathbf{y}$  is not one-to-one, the  $\mathbf{x}$ -space must be broken into regions where the function  $\mathbf{y}(\mathbf{x})$  can be inverted, and the contributions to  $g(\mathbf{y})$  from each region summed.

Given a set of functions  $\mathbf{y} = (y_1, \dots, y_m)$  with  $m < n$ , one can construct  $n - m$  additional independent functions, apply the procedure above, then integrate the resulting  $g(\mathbf{y})$  over the unwanted  $y_i$  to find the marginal distribution of those of interest.

For a one-to-one transformation of discrete random variables, the probability is obtained by simple substitution; no Jacobian is necessary because in this case  $f$  is a probability rather than a probability density. If the transformation is not one-to-one, then one must sum the probabilities for all values of the original variable that contribute to a given value of the transformed variable. If  $f$  depends on a set of parameters  $\boldsymbol{\theta}$ , a change to a different parameter set  $\boldsymbol{\eta}(\boldsymbol{\theta})$  is made by simple substitution; no Jacobian is used.

### 38.3. Characteristic functions

The characteristic function  $\phi(u)$  associated with the p.d.f.  $f(x)$  is essentially its Fourier transform, or the expectation value of  $e^{iux}$ :

$$\phi(u) = E[e^{iux}] = \int_{-\infty}^{\infty} e^{iux} f(x) dx. \quad (38.17)$$

Once  $\phi(u)$  is specified, the p.d.f.  $f(x)$  is uniquely determined and vice versa; knowing one is equivalent to the other. Characteristic functions are useful in deriving a number of important results about moments and sums of random variables.

It follows from Eqs. (38.8a) and (38.17) that the  $n^{\text{th}}$  moment of a random variable  $x$  that follows  $f(x)$  is given by

$$i^{-n} \left. \frac{d^n \phi}{du^n} \right|_{u=0} = \int_{-\infty}^{\infty} x^n f(x) dx = \alpha_n. \quad (38.18)$$

Thus it is often easy to calculate all the moments of a distribution defined by  $\phi(u)$ , even when  $f(x)$  cannot be written down explicitly.

If the p.d.f.s  $f_1(x)$  and  $f_2(y)$  for independent random variables  $x$  and  $y$  have characteristic functions  $\phi_1(u)$  and  $\phi_2(u)$ , then the characteristic function of the weighted sum  $ax + by$  is  $\phi_1(au)\phi_2(bu)$ . The rules of addition for several important distributions (*e.g.*, that the sum of two Gaussian distributed variables also follows a Gaussian distribution) easily follow from this observation.

Let the (partial) characteristic function corresponding to the conditional p.d.f.  $f_2(x|z)$  be  $\phi_2(u|z)$ , and the p.d.f. of  $z$  be  $f_1(z)$ . The characteristic function after integration over the conditional value is

$$\phi(u) = \int \phi_2(u|z)f_1(z) dz. \quad (38.19)$$

Suppose we can write  $\phi_2$  in the form

$$\phi_2(u|z) = A(u)e^{ig(u)z}. \quad (38.20)$$

Then

$$\phi(u) = A(u)\phi_1(g(u)). \quad (38.21)$$

The cumulants (semi-invariants)  $\kappa_n$  of a distribution with characteristic function  $\phi(u)$  are defined by the relation

$$\phi(u) = \exp \left[ \sum_{n=1}^{\infty} \frac{\kappa_n}{n!} (iu)^n \right] = \exp \left( i\kappa_1 u - \frac{1}{2}\kappa_2 u^2 + \dots \right). \quad (38.22)$$

The values  $\kappa_n$  are related to the moments  $\alpha_n$  and  $m_n$ . The first few relations are

$$\begin{aligned} \kappa_1 &= \alpha_1 (= \mu, \text{ the mean}) \\ \kappa_2 &= m_2 = \alpha_2 - \alpha_1^2 (= \sigma^2, \text{ the variance}) \\ \kappa_3 &= m_3 = \alpha_3 - 3\alpha_1\alpha_2 + 2\alpha_1^3. \end{aligned} \quad (38.23)$$

### 38.4. Commonly used probability distributions

Table 38.1 gives a number of common probability density functions and corresponding characteristic functions, means, and variances. Further information may be found in Refs. [1–8], [10], and [11], which has particularly detailed tables. Monte Carlo techniques for generating each of them may be found in our Sec. 40.4 and in Ref. [10]. We comment below on all except the trivial uniform distribution.

#### 38.4.1. Binomial and multinomial distributions :

A random process with exactly two possible outcomes which occur with fixed probabilities is called a *Bernoulli* process. If the probability of obtaining a certain outcome (a “success”) in an individual trial is  $p$ , then the probability of obtaining exactly  $r$  successes ( $r = 0, 1, 2, \dots, N$ ) in  $N$  independent trials, without regard to the order of the successes and failures, is given by the binomial distribution  $f(r; N, p)$  in Table 38.1. If  $r$  and  $s$  are binomially distributed with parameters  $(N_r, p)$  and  $(N_s, p)$ , then  $t = r + s$  follows a binomial distribution with parameters  $(N_r + N_s, p)$ .

If there are  $m$  possible outcomes for each trial having probabilities  $p_1, p_2, \dots, p_m$ , then the joint probability to find  $r_1, r_2, \dots, r_m$  of each outcome after a total of  $N$  independent trials is given by the multinomial distribution as shown in Table 38.1. We can regard outcome  $i$  as “success” and all the rest as “failure”, so individually, any of the  $r_i$  follow a binomial distribution for  $N$  trials and a success probability  $p_i$ .

**38.4.2. Poisson distribution :**

The Poisson distribution  $f(n; \nu)$  gives the probability of finding exactly  $n$  events in a given interval of  $x$  (e.g., space or time) when the events occur independently of one another and of  $x$  at an average rate of  $\nu$  per the given interval. The variance  $\sigma^2$  equals  $\nu$ . It is the limiting case  $p \rightarrow 0$ ,  $N \rightarrow \infty$ ,  $Np = \nu$  of the binomial distribution. The Poisson distribution approaches the Gaussian distribution for large  $\nu$ .

For example, a large number of radioactive nuclei of a given type will result in a certain number of decays in a fixed time interval. If this interval is small compared to the mean lifetime, then the probability for a given nucleus to decay is small, and thus the number of decays in the time interval is well modeled as a Poisson variable.

**38.4.3. Normal or Gaussian distribution :**

The normal (or Gaussian) probability density function  $f(x; \mu, \sigma^2)$  given in Table 38.1 has mean  $E[x] = \mu$  and variance  $V[x] = \sigma^2$ . Comparison of the characteristic function  $\phi(u)$  given in Table 38.1 with Eq. (38.22) shows that all cumulants  $\kappa_n$  beyond  $\kappa_2$  vanish; this is a unique property of the Gaussian distribution. Some other properties are:

$$P(x \text{ in range } \mu \pm \sigma) = 0.6827,$$

$$P(x \text{ in range } \mu \pm 0.6745\sigma) = 0.5,$$

$$E[|x - \mu|] = \sqrt{2/\pi}\sigma = 0.7979\sigma,$$

$$\text{half-width at half maximum} = \sqrt{2 \ln 2} \sigma = 1.177\sigma.$$

**Table 38.1.** Some common probability density functions, with corresponding characteristic functions and means and variances. In the Table,  $\Gamma(k)$  is the gamma function, equal to  $(k-1)!$  when  $k$  is an integer;  ${}_1F_1$  is the confluent hypergeometric function of the 1st kind [11].

Distribution	Probability density function $f$ (variable; parameters)	Characteristic function $\phi(u)$	Mean	Variance
Uniform	$f(x; a, b) = \begin{cases} 1/(b-a) & a \leq x \leq b \\ 0 & \text{otherwise} \end{cases}$	$\frac{e^{ibu} - e^{iau}}{(b-a)iu}$	$\frac{a+b}{2}$	$\frac{(b-a)^2}{12}$
Binomial	$f(r; N, p) = \frac{N!}{r!(N-r)!} p^r q^{N-r}$ $r = 0, 1, 2, \dots, N; \quad 0 \leq p \leq 1; \quad q = 1 - p$	$(q + pe^{iu})^N$	$Np$	$Npq$
Multinomial	$f(r_1, \dots, r_m; N, p_1, \dots, p_m) = \frac{N!}{r_1! \dots r_m!} p_1^{r_1} \dots p_m^{r_m}$ $r_k = 0, 1, 2, \dots, N; \quad 0 \leq p_k \leq 1; \quad \sum_{k=1}^m r_k = N$	$(\sum_{k=1}^m p_k e^{iu_k})^N$	$E[r_i] = Np_i$	$\text{cov}[r_i, r_j] = Np_i(\delta_{ij} - p_j)$
Poisson	$f(n; \nu) = \frac{\nu^n e^{-\nu}}{n!}; \quad n = 0, 1, 2, \dots; \quad \nu > 0$	$\exp[\nu(e^{iu} - 1)]$	$\nu$	$\nu$
Normal (Gaussian)	$f(x; \mu, \sigma^2) = \frac{1}{\sigma\sqrt{2\pi}} \exp(-(x-\mu)^2/2\sigma^2)$ $-\infty < x < \infty; \quad -\infty < \mu < \infty; \quad \sigma > 0$	$\exp(i\mu u - \frac{1}{2}\sigma^2 u^2)$	$\mu$	$\sigma^2$
Multivariate Gaussian	$f(\mathbf{x}; \boldsymbol{\mu}, V) = \frac{1}{(2\pi)^{n/2} \sqrt{ V }} \times \exp[-\frac{1}{2}(\mathbf{x} - \boldsymbol{\mu})^T V^{-1}(\mathbf{x} - \boldsymbol{\mu})]$ $-\infty < x_j < \infty; \quad -\infty < \mu_j < \infty; \quad  V  > 0$	$\exp[i\boldsymbol{\mu} \cdot \mathbf{u} - \frac{1}{2}\mathbf{u}^T V \mathbf{u}]$	$\boldsymbol{\mu}$	$V_{jk}$
Log-normal	$f(x; \mu, \sigma^2) = \frac{1}{\sigma\sqrt{2\pi}} \frac{1}{x} \exp(-(\ln x - \mu)^2/2\sigma^2)$ $0 < x < \infty; \quad -\infty < \mu < \infty; \quad \sigma > 0$	—	$\exp(\mu + \sigma^2/2)$	$\frac{\exp(2\mu + \sigma^2)}{\times [\exp(\sigma^2) - 1]}$
$\chi^2$	$f(z; n) = \frac{z^{n/2-1} e^{-z/2}}{2^{n/2} \Gamma(n/2)}; \quad z \geq 0$	$(1 - 2iu)^{-n/2}$	$n$	$2n$
Student's $t$	$f(t; n) = \frac{1}{\sqrt{n\pi}} \frac{\Gamma[(n+1)/2]}{\Gamma(n/2)} \left(1 + \frac{t^2}{n}\right)^{-(n+1)/2}$ $-\infty < t < \infty; \quad n \text{ not required to be integer}$	—	$0$ for $n > 1$	$n/(n-2)$ for $n > 2$
Gamma	$f(x; \lambda, k) = \frac{x^{k-1} \lambda^k e^{-\lambda x}}{\Gamma(k)}; \quad 0 \leq x < \infty;$ $k \text{ not required to be integer}$	$(1 - iu/\lambda)^{-k}$	$k/\lambda$	$k/\lambda^2$
Beta	$f(x; \alpha, \beta) = \frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} x^{\alpha-1} (1-x)^{\beta-1}$ $0 \leq x \leq 1$	${}_1F_1(\alpha; \alpha+\beta; iu)$	$\frac{\alpha}{\alpha+\beta}$	$\frac{\alpha\beta}{(\alpha+\beta)^2(\alpha+\beta+1)}$

For a Gaussian with  $\mu = 0$  and  $\sigma^2 = 1$  (the *standard normal*) the cumulative distribution, often written  $\Phi(x)$ , is related to the error function  $\text{erf}$  by

$$F(x; 0, 1) \equiv \Phi(x) = \frac{1}{2} \left[ 1 + \text{erf}(x/\sqrt{2}) \right]. \quad (38.24)$$

The error function and standard Gaussian are tabulated in many references (e.g., Ref. [11,12]) and are available in software packages such as ROOT [13]. For a mean  $\mu$  and variance  $\sigma^2$ , replace  $x$  by  $(x - \mu)/\sigma$ . The probability of  $x$  in a given range can be calculated with Eq. (39.65).

For  $x$  and  $y$  independent and normally distributed,  $z = ax + by$  follows a normal p.d.f.  $f(z; a\mu_x + b\mu_y, a^2\sigma_x^2 + b^2\sigma_y^2)$ ; that is, the weighted means and variances add.

The Gaussian derives its importance in large part from the *central limit theorem*:

If independent random variables  $x_1, \dots, x_n$  are distributed according to any p.d.f. with finite mean and variance, then the sum  $y = \sum_{i=1}^n x_i$  will have a p.d.f. that approaches a Gaussian for large  $n$ . If the p.d.f.s of the  $x_i$  are not identical, the theorem still holds under somewhat more restrictive conditions. The mean and variance are given by the sums of corresponding terms from the individual  $x_i$ . Therefore, the sum of a large number of fluctuations  $x_i$  will be distributed as a Gaussian, even if the  $x_i$  themselves are not.

For a set of  $n$  Gaussian random variables  $\mathbf{x}$  with means  $\boldsymbol{\mu}$  and covariances  $V_{ij} = \text{cov}[x_i, x_j]$ , the p.d.f. for the one-dimensional Gaussian is given by

$$f(\mathbf{x}; \boldsymbol{\mu}, V) = \frac{1}{(2\pi)^{n/2} \sqrt{|V|}} \exp \left[ -\frac{1}{2} (\mathbf{x} - \boldsymbol{\mu})^T V^{-1} (\mathbf{x} - \boldsymbol{\mu}) \right], \quad (38.25)$$

where the determinant  $|V|$  must be greater than 0. For diagonal  $V$  (independent variables),  $f(\mathbf{x}; \boldsymbol{\mu}, V)$  is the product of the p.d.f.s of  $n$  Gaussian distributions.

For  $n = 2$ ,  $f(\mathbf{x}; \boldsymbol{\mu}, V)$  is

$$f(x_1, x_2; \mu_1, \mu_2, \sigma_1, \sigma_2, \rho) = \frac{1}{2\pi\sigma_1\sigma_2\sqrt{1-\rho^2}} \times \exp \left\{ \frac{-1}{2(1-\rho^2)} \left[ \frac{(x_1 - \mu_1)^2}{\sigma_1^2} - \frac{2\rho(x_1 - \mu_1)(x_2 - \mu_2)}{\sigma_1\sigma_2} + \frac{(x_2 - \mu_2)^2}{\sigma_2^2} \right] \right\}. \quad (38.26)$$

The characteristic function for the multivariate Gaussian is

$$\phi(\mathbf{u}; \boldsymbol{\mu}, V) = \exp \left[ i\boldsymbol{\mu} \cdot \mathbf{u} - \frac{1}{2} \mathbf{u}^T V \mathbf{u} \right]. \quad (38.27)$$

If the components of  $\mathbf{x}$  are independent, then Eq. (38.27) is the product of the characteristic functions of  $n$  Gaussians.

For an  $n$ -dimensional Gaussian distribution for  $\mathbf{x}$  with mean  $\boldsymbol{\mu}$  and covariance matrix  $V$ , the marginal distribution for any single  $x_i$  is a one-dimensional Gaussian with mean  $\mu_i$  and variance  $V_{ii}$ . The equation  $(\mathbf{x} - \mathbf{a})^T V^{-1} (\mathbf{x} - \mathbf{a}) = C$ , where  $C$  is any positive number, defines an  $n$ -dimensional ellipse centered about  $\mathbf{a}$ . If  $\mathbf{a}$  is equal to the mean  $\boldsymbol{\mu}$ , then  $C$  is a random variable obeying the  $\chi^2$  distribution for  $n$  degrees of freedom, which is discussed in the following section. The probability that  $\mathbf{x}$  lies outside the ellipsoid for a given value of  $C$  is given by  $1 - F_{\chi^2}(C; n)$ , where  $F_{\chi^2}$  is the cumulative  $\chi^2$  distribution. This may be read from Fig. 39.1. For example, the “standard-deviation ellipsoid” occurs at  $C = s^2$ . For the two-variable case ( $n = 2$ ), the point  $\mathbf{x}$  lies outside the one-standard-deviation ellipsoid with 61% probability. The use of these ellipsoids as indicators of probable error is described in Sec. 39.4.2.2; the validity of those indicators assumes that  $\boldsymbol{\mu}$  and  $V$  are correct.

#### 38.4.4. Log-normal distribution :

If a random variable  $y$  follows a Gaussian distribution with mean  $\mu$  and variance  $\sigma^2$ , then  $x = e^y$  follows a log-normal distribution, as given in Table 38.1. As a consequence of the central limit theorem described in Sec. 38.4.3, the distribution of the product of a large number of positive random variables approaches a log-normal. It is bounded below by zero and is thus well suited for modeling quantities that are intrinsically non-negative such as an efficiency. One can implement a log-normal model for a random variable  $x$  by defining  $y = \ln x$  so that  $y$  follows a Gaussian distribution.

#### 38.4.5. $\chi^2$ distribution :

If  $x_1, \dots, x_n$  are independent Gaussian random variables, the sum  $z = \sum_{i=1}^n (x_i - \mu_i)^2 / \sigma_i^2$  follows the  $\chi^2$  p.d.f. with  $n$  degrees of freedom, which we denote by  $\chi^2(n)$ . More generally, for  $n$  correlated Gaussian variables as components of a vector  $\mathbf{X}$  with covariance matrix  $V$ ,  $z = \mathbf{X}^T V^{-1} \mathbf{X}$  follows  $\chi^2(n)$  as in the previous section. For a set of  $z_i$ , each of which follows  $\chi^2(n_i)$ ,  $\sum z_i$  follows  $\chi^2(\sum n_i)$ . For large  $n$ , the  $\chi^2$  p.d.f. approaches a Gaussian with a mean and variance given by  $\mu = n$  and  $\sigma^2 = 2n$ , respectively (here the formulae for  $\mu$  and  $\sigma^2$  are valid for all  $n$ ).

The  $\chi^2$  p.d.f. is often used in evaluating the level of compatibility between observed data and a hypothesis for the p.d.f. that the data might follow. This is discussed further in Sec. 39.3.2 on significance tests.

#### 38.4.6. Student's $t$ distribution :

Suppose that  $y$  and  $x_1, \dots, x_n$  are independent and Gaussian distributed with mean 0 and variance 1. We then define

$$z = \sum_{i=1}^n x_i^2 \quad \text{and} \quad t = \frac{y}{\sqrt{z/n}}. \quad (38.28)$$

The variable  $z$  thus follows a  $\chi^2(n)$  distribution. Then  $t$  is distributed according to Student's  $t$  distribution with  $n$  degrees of freedom,  $f(t; n)$ , given in Table 38.1.

If defined through gamma functions as in Table 38.1, the parameter  $n$  is not required to be an integer. As  $n \rightarrow \infty$ , the distribution approaches a Gaussian, and for  $n = 1$  it is a *Cauchy* or *Breit-Wigner* distribution.

As an example, consider the *sample mean*  $\bar{x} = \sum x_i / n$  and the *sample variance*  $s^2 = \sum (x_i - \bar{x})^2 / (n - 1)$  for normally distributed  $x_i$  with unknown mean  $\mu$  and variance  $\sigma^2$ . The sample mean has a Gaussian distribution with a variance  $\sigma^2/n$ , so the variable  $(\bar{x} - \mu) / \sqrt{\sigma^2/n}$  is normal with mean 0 and variance 1. The quantity  $(n - 1)s^2 / \sigma^2$  is independent of this and follows  $\chi^2(n - 1)$ . The ratio

$$t = \frac{(\bar{x} - \mu) / \sqrt{\sigma^2/n}}{\sqrt{(n - 1)s^2 / \sigma^2(n - 1)}} = \frac{\bar{x} - \mu}{\sqrt{s^2/n}} \quad (38.29)$$

is distributed as  $f(t; n - 1)$ . The unknown variance  $\sigma^2$  cancels, and  $t$  can be used to test the hypothesis that the true mean is some particular value  $\mu$ .

#### 38.4.7. Gamma distribution :

For a process that generates events as a function of  $x$  (e.g., space or time) according to a Poisson distribution, the distance in  $x$  from an arbitrary starting point (which may be some particular event) to the  $k^{\text{th}}$  event follows a *gamma* distribution,  $f(x; \lambda, k)$ . The Poisson parameter  $\mu$  is  $\lambda$  per unit  $x$ . The special case  $k = 1$  (i.e.,  $f(x; \lambda, 1) = \lambda e^{-\lambda x}$ ) is called the *exponential* distribution. A sum of  $k'$  exponential random variables  $x_i$  is distributed as  $f(\sum x_i; \lambda, k')$ .

The parameter  $k$  is not required to be an integer. For  $\lambda = 1/2$  and  $k = n/2$ , the gamma distribution reduces to the  $\chi^2(n)$  distribution.



**38.4.8. Beta distribution :**

The beta distribution describes a continuous random variable  $x$  in the interval  $[0, 1]$ . By scaling and translation one can easily generalize it to have arbitrary endpoints. In Bayesian inference about the parameter  $p$  of a binomial process, if the prior p.d.f. is a beta distribution  $f(p; \alpha, \beta)$  then the observation of  $r$  successes out of  $N$  trials gives a posterior beta distribution  $f(p; r + \alpha, N - r + \beta)$  (Bayesian methods are discussed further in Sec. 39). The uniform distribution is a beta distribution with  $\alpha = \beta = 1$ .

**References:**

1. H. Cramér, *Mathematical Methods of Statistics*, (Princeton Univ. Press, New Jersey, 1958).
2. A. Stuart and J.K. Ord, *Kendall's Advanced Theory of Statistics*, Vol. 1 *Distribution Theory* 6th Ed., (Halsted Press, New York, 1994), and earlier editions by Kendall and Stuart.
3. F.E. James, *Statistical Methods in Experimental Physics*, 2nd Ed., (World Scientific, Singapore, 2006).
4. L. Lyons, *Statistics for Nuclear and Particle Physicists*, (Cambridge University Press, New York, 1986).
5. B.R. Roe, *Probability and Statistics in Experimental Physics*, 2nd Ed., (Springer, New York, 2001).
6. R.J. Barlow, *Statistics: A Guide to the Use of Statistical Methods in the Physical Sciences*, (John Wiley, New York, 1989).
7. S. Brandt, *Data Analysis*, 3rd Ed., (Springer, New York, 1999).
8. G. Cowan, *Statistical Data Analysis*, (Oxford University Press, Oxford, 1998).
9. A.N. Kolmogorov, *Grundbegriffe der Wahrscheinlichkeitsrechnung*, (Springer, Berlin, 1933); *Foundations of the Theory of Probability*, 2nd Ed., (Chelsea, New York 1956).
10. Ch. Walck, *Hand-book on Statistical Distributions for Experimentalists*, University of Stockholm Internal Report SUF-PFY/96-01, available from [www.physto.se/~walck](http://www.physto.se/~walck).
11. M. Abramowitz and I. Stegun, eds., *Handbook of Mathematical Functions*, (Dover, New York, 1972).
12. F.W.J. Olver *et al.*, eds., *NIST Handbook of Mathematical Functions*, (Cambridge University Press, 2010); a companion Digital Library of Mathematical Functions is available at [dlmf.nist.gov](http://dlmf.nist.gov).
13. Rene Brun and Fons Rademakers, *Nucl. Instrum. Methods* **A389**, 81 (1997); see also [root.cern.ch](http://root.cern.ch).

## 39. STATISTICS

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This chapter gives an overview of statistical methods used in high-energy physics. In statistics, we are interested in using a given sample of data to make inferences about a probabilistic model, *e.g.*, to assess the model's validity or to determine the values of its parameters. There are two main approaches to statistical inference, which we may call frequentist and Bayesian.

In frequentist statistics, probability is interpreted as the frequency of the outcome of a repeatable experiment. The most important tools in this framework are parameter estimation, covered in Section 39.2, statistical tests, discussed in Section 39.3, and confidence intervals, which are constructed so as to cover the true value of a parameter with a specified probability, as described in Section 39.4.2. Note that in frequentist statistics one does not define a probability for a hypothesis or for the value of a parameter.

In Bayesian statistics, the interpretation of probability is more general and includes *degree of belief* (called subjective probability). One can then speak of a probability density function (p.d.f.) for a parameter, which expresses one's state of knowledge about where its true value lies. Bayesian methods provide a natural means to include additional information, which in general may be subjective; in fact they *require* prior probabilities for the hypotheses (or parameters) in question, *i.e.*, the degree of belief about the parameters' values before carrying out the measurement. Using Bayes' theorem (Eq. (38.4)), the prior degree of belief is updated by the data from the experiment. Bayesian methods for interval estimation are discussed in Sections 39.4.1 and 39.4.2.4.

For many inference problems, the frequentist and Bayesian approaches give similar numerical values, even though they answer different questions and are based on fundamentally different interpretations of probability. In some important cases, however, the two approaches may yield very different results. For a discussion of Bayesian vs. non-Bayesian methods, see references written by a statistician [1], by a physicist [2], or the detailed comparison in Ref. 3.

Following common usage in physics, the word “error” is often used in this chapter to mean “uncertainty.” More specifically it can indicate the size of an interval as in “the standard error” or “error propagation,” where the term refers to the standard deviation of an estimator.

## 39.1. Fundamental concepts

Consider an experiment whose outcome is characterized by one or more data values, which we can write as a vector  $\mathbf{x}$ . A *hypothesis*  $H$  is a statement about the probability for the data, often written  $P(\mathbf{x}|H)$ . (We will usually use a capital letter for a probability and lower case for a probability density. Often the term p.d.f. is used loosely to refer to either a probability or a probability density.) This could, for example, define completely the p.d.f. for the data (a simple hypothesis), or it could specify only the functional form of the p.d.f., with the values of one or more parameters not determined (a composite hypothesis).

If the probability  $P(\mathbf{x}|H)$  for data  $\mathbf{x}$  is regarded as a function of the hypothesis  $H$ , then it is called the *likelihood* of  $H$ , usually written  $L(H)$ . Often the hypothesis is characterized by one or more parameters  $\theta$ , in which case  $L(\theta) = P(\mathbf{x}|\theta)$  is called the likelihood function.

In some cases one can obtain at least approximate frequentist results using the likelihood evaluated only with the data obtained. In general, however, the frequentist approach requires a full specification of the probability model  $P(\mathbf{x}|H)$  both as a function of the data  $\mathbf{x}$  and hypothesis  $H$ .

In the Bayesian approach, inference is based on the posterior probability for  $H$  given the data  $\mathbf{x}$ , which represents one's degree of

belief that  $H$  is true given the data. This is obtained from Bayes' theorem (38.4), which can be written

$$P(H|\mathbf{x}) = \frac{P(\mathbf{x}|H)\pi(H)}{\int P(\mathbf{x}|H')\pi(H')dH'}. \quad (39.1)$$

Here  $P(\mathbf{x}|H)$  is the likelihood for  $H$ , which depends only on the data actually obtained. The quantity  $\pi(H)$  is the prior probability for  $H$ , which represents one's degree of belief for  $H$  before carrying out the measurement. The integral in the denominator (or sum, for discrete hypotheses) serves as a normalization factor. If  $H$  is characterized by a continuous parameter  $\theta$  then the posterior probability is a p.d.f.  $p(\theta|\mathbf{x})$ . Note that the likelihood function itself is not a p.d.f. for  $\theta$ .

## 39.2. Parameter estimation

Here we review *point estimation* of parameters, first with an overview of the frequentist approach and its two most important methods, maximum likelihood and least squares, treated in Sections 39.2.2 and 39.2.3. The Bayesian approach is outlined in Sec. 39.2.4.

An *estimator*  $\hat{\theta}$  (written with a hat) is a function of the data used to estimate the value of the parameter  $\theta$ . Sometimes the word ‘estimate’ is used to denote the value of the estimator when evaluated with given data. There is no fundamental rule dictating how an estimator must be constructed. One tries, therefore, to choose that estimator which has the best properties. The most important of these are (a) *consistency*, (b) *bias*, (c) *efficiency*, and (d) *robustness*.

(a) An estimator is said to be *consistent* if the estimate  $\hat{\theta}$  converges in probability (see Ref. 3) to the true value  $\theta$  as the amount of data increases. This property is so important that it is possessed by all commonly used estimators.

(b) The *bias*,  $b = E[\hat{\theta}] - \theta$ , is the difference between the expectation value of the estimator and the true value of the parameter. The expectation value is taken over a hypothetical set of similar experiments in which  $\hat{\theta}$  is constructed in the same way. When  $b = 0$ , the estimator is said to be unbiased. The bias depends on the chosen metric, *i.e.*, if  $\hat{\theta}$  is an unbiased estimator of  $\theta$ , then  $\hat{\theta}^2$  is not in general an unbiased estimator for  $\theta^2$ .

(c) *Efficiency* is the ratio of the minimum possible variance for any estimator of  $\theta$  to the variance  $V[\hat{\theta}]$  of the estimator  $\hat{\theta}$ . For the case of a single parameter, under rather general conditions the minimum variance is given by the Rao-Cramér-Fréchet bound,

$$\sigma_{\min}^2 = \left(1 + \frac{\partial b}{\partial \theta}\right)^2 / I(\theta), \quad (39.2)$$

where

$$I(\theta) = E \left[ \left( \frac{\partial \ln L}{\partial \theta} \right)^2 \right] = -E \left[ \frac{\partial^2 \ln L}{\partial \theta^2} \right] \quad (39.3)$$

is the *Fisher information*,  $L$  is the likelihood, and the expectation value in (39.3) is carried out with respect to the data. For the final equality to hold, the range of allowed data values must not depend on  $\theta$ .

The *mean-squared error*,

$$\text{MSE} = E[(\hat{\theta} - \theta)^2] = V[\hat{\theta}] + b^2, \quad (39.4)$$

is a measure of an estimator's quality which combines bias and variance.

(d) *Robustness* is the property of being insensitive to departures from assumptions in the p.d.f., *e.g.*, owing to uncertainties in the distribution's tails.

It is not in general possible to optimize simultaneously for all the measures of estimator quality described above. For example, there is in general a trade-off between bias and variance. For some common estimators, the properties above are known exactly. More generally, it is possible to evaluate them by Monte Carlo simulation. Note that they will in general depend on the unknown  $\theta$ .

### 39.2.1. Estimators for mean, variance, and median :

Suppose we have a set of  $n$  independent measurements,  $x_1, \dots, x_n$ , each assumed to follow a p.d.f. with unknown mean  $\mu$  and unknown variance  $\sigma^2$  (the measurements do not necessarily have to follow a Gaussian distribution). Then

$$\hat{\mu} = \frac{1}{n} \sum_{i=1}^n x_i \quad (39.5)$$

$$\hat{\sigma}^2 = \frac{1}{n-1} \sum_{i=1}^n (x_i - \hat{\mu})^2 \quad (39.6)$$

are unbiased estimators of  $\mu$  and  $\sigma^2$ . The variance of  $\hat{\mu}$  is  $\sigma^2/n$  and the variance of  $\hat{\sigma}^2$  is

$$V[\hat{\sigma}^2] = \frac{1}{n} \left( m_4 - \frac{n-3}{n-1} \sigma^4 \right), \quad (39.7)$$

where  $m_4$  is the 4th central moment of  $x$  (see Eq. (38.8b)). For Gaussian distributed  $x_i$ , this becomes  $2\sigma^4/(n-1)$  for any  $n \geq 2$ , and for large  $n$  the standard deviation of  $\hat{\sigma}$  (the “error of the error”) is  $\sigma/\sqrt{2n}$ . For any  $n$  and Gaussian  $x_i$ ,  $\hat{\mu}$  is an efficient estimator for  $\mu$ , and the estimators  $\hat{\mu}$  and  $\hat{\sigma}^2$  are uncorrelated. Otherwise the arithmetic mean (39.5) is not necessarily the most efficient estimator; this is discussed further in Sec. 8.7 of Ref. 4.

If  $\sigma^2$  is known, it does not improve the estimate  $\hat{\mu}$ , as can be seen from Eq. (39.5); however, if  $\mu$  is known, one can substitute it for  $\hat{\mu}$  in Eq. (39.6) and replace  $n-1$  by  $n$  to obtain an estimator of  $\sigma^2$  still with zero bias but smaller variance. If the  $x_i$  have different, known variances  $\sigma_i^2$ , then the weighted average

$$\hat{\mu} = \frac{1}{w} \sum_{i=1}^n w_i x_i, \quad (39.8)$$

where  $w_i = 1/\sigma_i^2$  and  $w = \sum_i w_i$ , is an unbiased estimator for  $\mu$  with a smaller variance than an unweighted average. The standard deviation of  $\hat{\mu}$  is  $1/\sqrt{w}$ .

As an estimator for the median  $x_{\text{med}}$ , one can use the value  $\hat{x}_{\text{med}}$  such that half the  $x_i$  are below and half above (the sample median). If there are an even number of observations and the sample median lies between two observed values, the estimator is set by convention to their arithmetic average. If the p.d.f. of  $x$  has the form  $f(x - \mu)$  and  $\mu$  is both mean and median, then for large  $n$  the variance of the sample median approaches  $1/[4nf^2(0)]$ , provided  $f(0) > 0$ . Although estimating the median can often be more difficult computationally than the mean, the resulting estimator is generally more robust, as it is insensitive to the exact shape of the tails of a distribution.

### 39.2.2. The method of maximum likelihood :

Suppose we have a set of measured quantities  $\mathbf{x}$  and the likelihood  $L(\boldsymbol{\theta}) = P(\mathbf{x}|\boldsymbol{\theta})$  for a set of parameters  $\boldsymbol{\theta} = (\theta_1, \dots, \theta_N)$ . The *maximum likelihood* (ML) estimators for  $\boldsymbol{\theta}$  are defined as the values that give the maximum of  $L$ . Because of the properties of the logarithm, it is usually easier to work with  $\ln L$ , and since both are maximized for the same parameter values  $\boldsymbol{\theta}$ , the ML estimators can be found by solving the *likelihood equations*,

$$\frac{\partial \ln L}{\partial \theta_i} = 0, \quad i = 1, \dots, N. \quad (39.9)$$

Often the solution must be found numerically. Maximum likelihood estimators are important because they are unbiased and efficient asymptotically (i.e., for large data samples), under quite general conditions, and the method has a wide range of applicability.

In general the likelihood function is obtained from the probability of the data under assumption of the parameters. An important special case is when the data consist of *i.i.d.* (independent and identically distributed) values. Here one has a set of  $n$  statistically independent quantities  $\mathbf{x} = (x_1, \dots, x_n)$ , where each component follows the same

p.d.f.  $f(x; \boldsymbol{\theta})$ . In this case the joint p.d.f. of the data sample factorizes and the likelihood function is

$$L(\boldsymbol{\theta}) = \prod_{i=1}^n f(x_i; \boldsymbol{\theta}). \quad (39.10)$$

In this case the number of events  $n$  is regarded as fixed. If however the probability to observe  $n$  events itself depends on the parameters  $\boldsymbol{\theta}$ , then this dependence should be included in the likelihood. For example, if  $n$  follows a Poisson distribution with mean  $\mu$  and the independent  $x$  values all follow  $f(x; \boldsymbol{\theta})$ , then the likelihood becomes

$$L(\boldsymbol{\theta}) = \frac{\mu^n}{n!} e^{-\mu} \prod_{i=1}^n f(x_i; \boldsymbol{\theta}). \quad (39.11)$$

Equation (39.11) is often called the *extended likelihood* (see, e.g., Refs. [6–8]). If  $\mu$  is given as a function of  $\boldsymbol{\theta}$ , then including the probability for  $n$  given  $\boldsymbol{\theta}$  in the likelihood provides additional information about the parameters. This therefore leads to a reduction in their statistical uncertainties and in general changes their estimated values.

In evaluating the likelihood function, it is important that any normalization factors in the p.d.f. that involve  $\boldsymbol{\theta}$  be included. However, we will only be interested in the maximum of  $L$  and in ratios of  $L$  at different values of the parameters; hence any multiplicative factors that do not involve the parameters that we want to estimate may be dropped, including factors that depend on the data but not on  $\boldsymbol{\theta}$ .

Under a one-to-one change of parameters from  $\boldsymbol{\theta}$  to  $\boldsymbol{\eta}$ , the ML estimators  $\hat{\boldsymbol{\theta}}$  transform to  $\boldsymbol{\eta}(\hat{\boldsymbol{\theta}})$ . That is, the ML solution is invariant under change of parameter. However, other properties of ML estimators, in particular the bias, are not invariant under change of parameter.

The inverse  $V^{-1}$  of the covariance matrix  $V_{ij} = \text{cov}[\hat{\theta}_i, \hat{\theta}_j]$  for a set of ML estimators can be estimated by using

$$(\hat{V}^{-1})_{ij} = - \frac{\partial^2 \ln L}{\partial \theta_i \partial \theta_j} \bigg|_{\hat{\boldsymbol{\theta}}}; \quad (39.12)$$

for finite samples, however, Eq. (39.12) can result in a misestimate of the variances. In the large sample limit (or in a linear model with Gaussian errors),  $L$  has a Gaussian form and  $\ln L$  is (hyper)parabolic. In this case, it can be seen that a numerically equivalent way of determining  $s$ -standard-deviation errors is from the hypersurface defined by the  $\boldsymbol{\theta}$  such that

$$\ln L(\boldsymbol{\theta}) = \ln L_{\text{max}} - s^2/2, \quad (39.13)$$

where  $\ln L_{\text{max}}$  is the value of  $\ln L$  at the solution point (compare with Eq. (39.68)). The minimum and maximum values of  $\theta_i$  on the hypersurface then give an approximate  $s$ -standard deviation confidence interval for  $\theta_i$  (see Section 39.4.2.2).

#### 39.2.2.1. ML with binned data:

If the total number of data values  $x_i$ ,  $i = 1, \dots, n_{\text{tot}}$ , is small, the unbinned maximum likelihood method, i.e., use of Equation (39.10) (or (39.11) for extended ML), is preferred since binning can only result in a loss of information, and hence larger statistical errors for the parameter estimates. If the sample is large, it can be convenient to bin the values in a histogram with  $N$  bins, so that one obtains a vector of data  $\mathbf{n} = (n_1, \dots, n_N)$  with expectation values  $\boldsymbol{\mu} = E[\mathbf{n}]$  and probabilities  $f(\mathbf{n}; \boldsymbol{\mu})$ . Suppose the mean values  $\boldsymbol{\mu}$  can be determined as a function of a set of parameters  $\boldsymbol{\theta}$ . Then one may maximize the likelihood function based on the contents of the bins.

As mentioned in Sec. 39.2.2, the total number of events  $n_{\text{tot}} = \sum_i n_i$  can be regarded either as fixed or as a random variable. If it is fixed, the histogram follows a multinomial distribution,

$$f_M(\mathbf{n}; \boldsymbol{\theta}) = \frac{n_{\text{tot}}!}{n_1! \dots n_N!} p_1^{n_1} \dots p_N^{n_N}, \quad (39.14)$$

where we assume the probabilities  $p_i$  are given functions of the parameters  $\boldsymbol{\theta}$ . The distribution can be written equivalently in terms

of the expected number of events in each bin,  $\mu_i = n_{\text{tot}} p_i$ . If the  $n_i$  are regarded as independent and Poisson distributed, then the data are instead described by a product of Poisson probabilities,

$$f_P(\mathbf{n}; \boldsymbol{\theta}) = \prod_{i=1}^N \frac{\mu_i^{n_i}}{n_i!} e^{-\mu_i}, \quad (39.15)$$

where the mean values  $\mu_i$  are given functions of  $\boldsymbol{\theta}$ . The total number of events  $n_{\text{tot}}$  thus follows a Poisson distribution with mean  $\mu_{\text{tot}} = \sum_i \mu_i$ .

When using maximum likelihood with binned data, one can find the ML estimators and at the same time obtain a statistic usable for a test of goodness-of-fit (see Sec. 39.3.2). Maximizing the likelihood  $L(\boldsymbol{\theta}) = f_{M/P}(\mathbf{n}; \boldsymbol{\theta})$  is equivalent to maximizing the likelihood ratio  $\lambda(\boldsymbol{\theta}) = f_{M/P}(\mathbf{n}; \boldsymbol{\theta}) / f(\mathbf{n}; \hat{\boldsymbol{\mu}})$ , where in the denominator  $f(\mathbf{n}; \boldsymbol{\mu})$  is a model with an adjustable parameter for each bin,  $\boldsymbol{\mu} = (\mu_1, \dots, \mu_N)$ , and the corresponding estimators are  $\hat{\boldsymbol{\mu}} = (n_1, \dots, n_N)$ . Equivalently one often minimizes the quantity  $-2 \ln \lambda(\boldsymbol{\theta})$ . For independent Poisson distributed  $n_i$  this is [9]

$$-2 \ln \lambda(\boldsymbol{\theta}) = 2 \sum_{i=1}^N \left[ \mu_i(\boldsymbol{\theta}) - n_i + n_i \ln \frac{n_i}{\mu_i(\boldsymbol{\theta})} \right], \quad (39.16)$$

where for bins with  $n_i = 0$ , the last term in (39.16) is zero. The expression (39.16) without the terms  $\mu_i - n_i$  also gives  $-2 \ln \lambda(\boldsymbol{\theta})$  for multinomially distributed  $n_i$ , *i.e.*, when the total number of entries is regarded as fixed. In the limit of zero bin width, minimizing (39.16) is equivalent to maximizing the unbinned extended likelihood function (39.11) or in the multinomial case without the  $\mu_i - n_i$  terms one obtains Eq. (39.10).

A smaller value of  $-2 \ln \lambda(\hat{\boldsymbol{\theta}})$  corresponds to better agreement between the data and the hypothesized form of  $\boldsymbol{\mu}(\boldsymbol{\theta})$ . The value of  $-2 \ln \lambda(\hat{\boldsymbol{\theta}})$  can thus be translated into a  $p$ -value as a measure of goodness-of-fit, as described in Sec. 39.3.2. Assuming the model is correct, then according to Wilks' theorem [10], for sufficiently large  $\mu_i$  and provided certain regularity conditions are met, the minimum of  $-2 \ln \lambda$  as defined by Eq. (39.16) follows a  $\chi^2$  distribution (see, *e.g.*, Ref. 9). If there are  $N$  bins and  $m$  fitted parameters, then the number of degrees of freedom for the  $\chi^2$  distribution is  $N - m$  if the data are treated as Poisson-distributed, and  $N - m - 1$  if the  $n_i$  are multinomially distributed.

Suppose the  $n_i$  are Poisson-distributed and the overall normalization  $\mu_{\text{tot}} = \sum_i \mu_i$  is taken as an adjustable parameter, so that  $\mu_i = \mu_{\text{tot}} p_i(\boldsymbol{\theta})$ , where the probability to be in the  $i$ th bin,  $p_i(\boldsymbol{\theta})$ , does not depend on  $\mu_{\text{tot}}$ . Then by minimizing Eq. (39.16), one obtains that the area under the fitted function is equal to the sum of the histogram contents, *i.e.*,  $\sum_i \hat{\mu}_i = \sum_i n_i$ . This is a property not possessed by the estimators from the method of least squares (see, *e.g.*, Sec. 39.2.3 and Ref. 8).

### 39.2.2.2. Frequentist treatment of nuisance parameters:

Suppose we want to determine the values of parameters  $\boldsymbol{\theta}$  using a set of measurements  $\mathbf{x}$  described by a probability model  $P_x(\mathbf{x}|\boldsymbol{\theta})$ . In general the model is not perfect, which is to say it can not provide an accurate description of the data even at the most optimal point of its parameter space. As a result, the estimated parameters can have a systematic bias.

One can improve the model by including in it additional parameters. That is,  $P_x(\mathbf{x}|\boldsymbol{\theta})$  is replaced by a more general model  $P_x(\mathbf{x}|\boldsymbol{\theta}, \boldsymbol{\nu})$ , which depends on parameters of interest  $\boldsymbol{\theta}$  and *nuisance parameters*  $\boldsymbol{\nu}$ . The additional parameters are not of intrinsic interest but must be included for the model to be accurate for some point in the enlarged parameter space.

Although including additional parameters may eliminate or at least reduce the effect of systematic uncertainties, their presence will result in increased statistical uncertainties for the parameters of interest. This occurs because the estimators for the nuisance parameters and those of interest will in general be correlated, which results in an enlargement of the contour defined by Eq. (39.13).

To reduce the impact of the nuisance parameters one often tries to constrain their values by means of control or calibration measurements, say, having data  $\mathbf{y}$ . For example, some components of  $\mathbf{y}$  could represent estimates of the nuisance parameters, often from separate experiments. Suppose the measurements  $\mathbf{y}$  are statistically independent from  $\mathbf{x}$  and are described by a model  $P_y(\mathbf{y}|\boldsymbol{\nu})$ . The joint model for both  $\mathbf{x}$  and  $\mathbf{y}$  is in this case therefore the product of the probabilities for  $\mathbf{x}$  and  $\mathbf{y}$ , and thus the likelihood function for the full set of parameters is

$$L(\boldsymbol{\theta}, \boldsymbol{\nu}) = P_x(\mathbf{x}|\boldsymbol{\theta}, \boldsymbol{\nu}) P_y(\mathbf{y}|\boldsymbol{\nu}). \quad (39.17)$$

Note that in this case if one wants to simulate the experiment by means of Monte Carlo, both the primary and control measurements,  $\mathbf{x}$  and  $\mathbf{y}$ , must be generated for each repetition under assumption of fixed values for the parameters  $\boldsymbol{\theta}$  and  $\boldsymbol{\nu}$ .

Using all of the parameters  $(\boldsymbol{\theta}, \boldsymbol{\nu})$  in Eq. (39.13) to find the statistical errors in the parameters of interest  $\boldsymbol{\theta}$  is equivalent to using the *profile likelihood*, which depends only on  $\boldsymbol{\theta}$ . It is defined as

$$L_p(\boldsymbol{\theta}) = L(\boldsymbol{\theta}, \hat{\boldsymbol{\nu}}(\boldsymbol{\theta})), \quad (39.18)$$

where the double-hat notation indicates the profiled values of the parameters  $\boldsymbol{\nu}$ , defined as the values that maximize  $L$  for the specified  $\boldsymbol{\theta}$ . The profile likelihood is discussed further in Section 39.3.2.1 in connection with hypothesis tests.

### 39.2.3. The method of least squares :

The *method of least squares* (LS) coincides with the method of maximum likelihood in the following special case. Consider a set of  $N$  independent measurements  $y_i$  at known points  $x_i$ . The measurement  $y_i$  is assumed to be Gaussian distributed with mean  $\mu(x_i; \boldsymbol{\theta})$  and known variance  $\sigma_i^2$ . The goal is to construct estimators for the unknown parameters  $\boldsymbol{\theta}$ . The log-likelihood function contains the sum of squares

$$\chi^2(\boldsymbol{\theta}) = -2 \ln L(\boldsymbol{\theta}) + \text{constant} = \sum_{i=1}^N \frac{(y_i - \mu(x_i; \boldsymbol{\theta}))^2}{\sigma_i^2}. \quad (39.19)$$

The parameter values that maximize  $L$  are the same as those which minimize  $\chi^2$ .

The minimum of the chi-square function in Equation (39.19) defines the least-squares estimators  $\hat{\boldsymbol{\theta}}$  for the more general case where the  $y_i$  are not Gaussian distributed as long as they are independent. If they are not independent but rather have a covariance matrix  $V_{ij} = \text{cov}[y_i, y_j]$ , then the LS estimators are determined by the minimum of

$$\chi^2(\boldsymbol{\theta}) = (\mathbf{y} - \boldsymbol{\mu}(\boldsymbol{\theta}))^T V^{-1} (\mathbf{y} - \boldsymbol{\mu}(\boldsymbol{\theta})), \quad (39.20)$$

where  $\mathbf{y} = (y_1, \dots, y_N)$  is the (column) vector of measurements,  $\boldsymbol{\mu}(\boldsymbol{\theta})$  is the corresponding vector of predicted values, and the superscript  $T$  denotes the transpose. If the  $y_i$  are not Gaussian distributed, then the LS and ML estimators will not in general coincide.

Often one further restricts the problem to the case where  $\mu(x_i; \boldsymbol{\theta})$  is a linear function of the parameters, *i.e.*,

$$\mu(x_i; \boldsymbol{\theta}) = \sum_{j=1}^m \theta_j h_j(x_i). \quad (39.21)$$

Here the  $h_j(x)$  are  $m$  linearly independent functions, *e.g.*,  $1, x, x^2, \dots, x^{m-1}$  or Legendre polynomials. We require  $m < N$  and at least  $m$  of the  $x_i$  must be distinct.

Minimizing  $\chi^2$  in this case with  $m$  parameters reduces to solving a system of  $m$  linear equations. Defining  $H_{ij} = h_j(x_i)$  and minimizing  $\chi^2$  by setting its derivatives with respect to the  $\theta_i$  equal to zero gives the LS estimators,

$$\hat{\boldsymbol{\theta}} = (H^T V^{-1} H)^{-1} H^T V^{-1} \mathbf{y} \equiv D \mathbf{y}. \quad (39.22)$$

The covariance matrix for the estimators  $U_{ij} = \text{cov}[\hat{\theta}_i, \hat{\theta}_j]$  is given by

$$U = DV D^T = (H^T V^{-1} H)^{-1}, \quad (39.23)$$

or equivalently, its inverse  $U^{-1}$  can be found from

$$(U^{-1})_{ij} = \frac{1}{2} \frac{\partial^2 \chi^2}{\partial \theta_i \partial \theta_j} \Big|_{\theta=\hat{\theta}} = \sum_{k,l=1}^N h_i(x_k) (V^{-1})_{kl} h_j(x_l). \quad (39.24)$$

The LS estimators can also be found from the expression

$$\hat{\theta} = U g, \quad (39.25)$$

where the vector  $g$  is defined by

$$g_i = \sum_{j,k=1}^N y_j h_i(x_k) (V^{-1})_{jk}. \quad (39.26)$$

For the case of uncorrelated  $y_i$ , for example, one can use (39.25) with

$$(U^{-1})_{ij} = \sum_{k=1}^N \frac{h_i(x_k) h_j(x_k)}{\sigma_k^2}, \quad (39.27)$$

$$g_i = \sum_{k=1}^N \frac{y_k h_i(x_k)}{\sigma_k^2}. \quad (39.28)$$

Expanding  $\chi^2(\theta)$  about  $\hat{\theta}$ , one finds that the contour in parameter space defined by

$$\chi^2(\theta) = \chi^2(\hat{\theta}) + 1 = \chi_{\min}^2 + 1 \quad (39.29)$$

has tangent planes located at approximately plus-or-minus-one standard deviation  $\sigma_{\hat{\theta}}$  from the LS estimates  $\hat{\theta}$ .

In constructing the quantity  $\chi^2(\theta)$  one requires the variances or, in the case of correlated measurements, the covariance matrix. Often these quantities are not known *a priori* and must be estimated from the data; an important example is where the measured value  $y_i$  represents the event count in a histogram bin. If, for example,  $y_i$  represents a Poisson variable, for which the variance is equal to the mean, then one can either estimate the variance from the predicted value,  $\mu(x_i; \theta)$ , or from the observed number itself,  $y_i$ . In the first option, the variances become functions of the parameters, and as a result the estimators may need to be found numerically. The second option can be undefined if  $y_i$  is zero, and for small  $y_i$ , the variance will be poorly estimated. In either case, one should constrain the normalization of the fitted curve to the correct value, *i.e.*, one should determine the area under the fitted curve directly from the number of entries in the histogram (see Ref. 8, Section 7.4). As noted in Sec. 39.2.2.1, this issue is avoided when using the method of extended maximum likelihood with binned data by minimizing Eq. (39.16). In that case if the expected number of events  $\mu_{\text{tot}}$  does not depend on the other fitted parameters  $\theta$ , then its extended ML estimator is equal to the observed total number of events.

As the minimum value of the  $\chi^2$  represents the level of agreement between the measurements and the fitted function, it can be used for assessing the goodness-of-fit; this is discussed further in Section 39.3.2.

### 39.2.4. The Bayesian approach :

In the frequentist methods discussed above, probability is associated only with data, not with the value of a parameter. This is no longer the case in Bayesian statistics, however, which we introduce in this section. For general introductions to Bayesian statistics see, *e.g.*, Refs. [24–27].

Suppose the outcome of an experiment is characterized by a vector of data  $\mathbf{x}$ , whose probability distribution depends on an unknown parameter (or parameters)  $\theta$  that we wish to determine. In Bayesian statistics, all knowledge about  $\theta$  is summarized by the posterior p.d.f.  $p(\theta|\mathbf{x})$ , whose integral over any given region gives the degree of belief

for  $\theta$  to take on values in that region, given the data  $\mathbf{x}$ . It is obtained by using Bayes' theorem,

$$p(\theta|\mathbf{x}) = \frac{P(\mathbf{x}|\theta)\pi(\theta)}{\int P(\mathbf{x}|\theta')\pi(\theta') d\theta'}, \quad (39.30)$$

where  $P(\mathbf{x}|\theta)$  is the likelihood function, *i.e.*, the joint p.d.f. for the data viewed as a function of  $\theta$ , evaluated with the data actually obtained in the experiment, and  $\pi(\theta)$  is the prior p.d.f. for  $\theta$ . Note that the denominator in Eq. (39.30) serves to normalize the posterior p.d.f. to unity.

As it can be difficult to report the full posterior p.d.f.  $p(\theta|\mathbf{x})$ , one would usually summarize it with statistics such as the mean (or median) value, and covariance matrix. In addition one may construct intervals with a given probability content, as is discussed in Sec. 39.4.1 on Bayesian interval estimation.

#### 39.2.4.1. Priors:

Bayesian statistics supplies no unique rule for determining the prior  $\pi(\theta)$ ; this reflects the analyst's subjective degree of belief (or state of knowledge) about  $\theta$  before the measurement was carried out. For the result to be of value to the broader community, whose members may not share these beliefs, it is important to carry out a *sensitivity analysis*, that is, to show how the result changes under a reasonable variation of the prior probabilities.

One might like to construct  $\pi(\theta)$  to represent complete ignorance about the parameters by setting it equal to a constant. A problem here is that if the prior p.d.f. is flat in  $\theta$ , then it is not flat for a nonlinear function of  $\theta$ , and so a different parametrization of the problem would lead in general to a non-equivalent posterior p.d.f.

For the special case of a constant prior, one can see from Bayes' theorem (39.30) that the posterior is proportional to the likelihood, and therefore the mode (peak position) of the posterior is equal to the ML estimator. The posterior mode, however, will change in general upon a transformation of parameter. One may use as the Bayesian estimator a summary statistic other than the mode, such as the median, which is invariant under parameter transformation. But this will not in general coincide with the ML estimator.

The difficult and subjective nature of encoding personal knowledge into priors has led to what is called *objective Bayesian statistics*, where prior probabilities are based not on an actual degree of belief but rather derived from formal rules. These give, for example, priors which are invariant under a transformation of parameters, or ones which result in a maximum gain in information for a given set of measurements. For an extensive review see, *e.g.*, Ref. 28.

Objective priors do not in general reflect degree of belief, but they could in some cases be taken as possible, although perhaps extreme, subjective priors. The posterior probabilities as well therefore do not necessarily reflect a degree of belief. However one may regard investigating a variety of objective priors to be an important part of the sensitivity analysis. Furthermore, use of objective priors with Bayes' theorem can be viewed as a recipe for producing estimators or intervals which have desirable frequentist properties.

An important procedure for deriving objective priors is due to Jeffreys. According to *Jeffreys' rule* one takes the prior as

$$\pi(\theta) \propto \sqrt{\det(I(\theta))}, \quad (39.31)$$

where

$$I_{ij}(\theta) = -E \left[ \frac{\partial^2 \ln P(\mathbf{x}|\theta)}{\partial \theta_i \partial \theta_j} \right] \quad (39.32)$$

is the *Fisher information matrix*. One can show that the Jeffreys prior leads to inference that is invariant under a transformation of parameters. One should note that the Jeffreys prior does not in general correspond to one's degree of belief about the value of a parameter. As examples, the Jeffreys prior for the mean  $\mu$  of a Gaussian distribution is a constant, and for the mean of a Poisson distribution one finds  $\pi(\mu) \propto 1/\sqrt{\mu}$ .

Neither the constant nor  $1/\sqrt{\mu}$  priors can be normalized to unit area and are therefore said to be *improper*. This can be allowed

because the prior always appears multiplied by the likelihood function, and if the likelihood falls to zero sufficiently quickly then one may have a normalizable posterior density.

An important type of objective prior is the reference prior due to Bernardo and Berger [29]. To find the reference prior for a given problem one considers the Kullback-Leibler divergence  $D_n[\pi, p]$  of the posterior  $p(\theta|\mathbf{x})$  relative to a prior  $\pi(\theta)$ , obtained from a set of i.i.d. data  $\mathbf{x} = (x_1, \dots, x_n)$ :

$$D_n[\pi, p] = \int p(\theta|\mathbf{x}) \ln \frac{p(\theta|\mathbf{x})}{\pi(\theta)} d\theta. \quad (39.33)$$

This is effectively a measure of the gain in information provided by the data. The reference prior is chosen so that the expectation value of this information gain is maximized for the limiting case of  $n \rightarrow \infty$ , where the expectation is computed with respect to the marginal distribution of the data,

$$p(\mathbf{x}) = \int p(\mathbf{x}|\theta)\pi(\theta) d\theta. \quad (39.34)$$

For a single, continuous parameter the reference prior is usually identical to the Jeffreys prior. In the multiparameter case an iterative algorithm exists, which requires sorting the parameters by order of inferential importance. Often the result does not depend on this order, but when it does, this can be part of a sensitivity analysis. Further discussion and applications to particle physics problems can be found in Ref. 30.

#### 39.2.4.2. Bayesian treatment of nuisance parameters:

As discussed in Sec. 39.2.2, a model may depend on parameters of interest  $\theta$  as well as on nuisance parameters  $\nu$ , which must be included for an accurate description of the data. Knowledge about the values of  $\nu$  may be supplied by control measurements, theoretical insights, physical constraints, etc. Suppose, for example, one has data  $\mathbf{y}$  from a control measurement which is characterized by a probability  $P_{\mathbf{y}}(\mathbf{y}|\nu)$ . Suppose further that before carrying out the control measurement one's state of knowledge about  $\nu$  is described by an initial prior  $\pi_0(\nu)$ , which in practice is often taken to be a constant or in any case very broad. By using Bayes' theorem (39.1) one obtains the updated prior  $\pi(\nu)$  (i.e., now  $\pi(\nu) = \pi(\nu|\mathbf{y})$ , the probability for  $\nu$  given  $\mathbf{y}$ ),

$$\pi(\nu|\mathbf{y}) \propto P(\mathbf{y}|\nu)\pi_0(\nu). \quad (39.35)$$

In the absence of a model for  $P(\mathbf{y}|\nu)$  one may make some reasonable but ad hoc choices. For a single nuisance parameter  $\nu$ , for example, one might characterize the uncertainty by a p.d.f.  $\pi(\nu)$  centered about its nominal value with a certain standard deviation  $\sigma_\nu$ . Often a Gaussian p.d.f. provides a reasonable model for one's degree of belief about a nuisance parameter; in other cases, more complicated shapes may be appropriate. If, for example, the parameter represents a non-negative quantity then a log-normal or gamma p.d.f. can be a more natural choice than a Gaussian truncated at zero. Note also that truncation of the prior of a nuisance parameter  $\nu$  at zero will in general make  $\pi(\nu)$  nonzero at  $\nu = 0$ , which can lead to an unnormalizable posterior for a parameter of interest that appears multiplied by  $\nu$ .

The likelihood function, prior, and posterior p.d.f.s all depend on both  $\theta$  and  $\nu$ , and are related by Bayes' theorem, as usual. Note that the likelihood here only refers to the primary measurement  $\mathbf{x}$ . Once any control measurements  $\mathbf{y}$  are used to find the updated prior  $\pi(\nu)$  for the nuisance parameters, this information is fully encapsulated in  $\pi(\nu)$  and the control measurements do not appear further.

One can obtain the posterior p.d.f. for  $\theta$  alone by integrating over the nuisance parameters, i.e.,

$$p(\theta|\mathbf{x}) = \int p(\theta, \nu|\mathbf{x}) d\nu. \quad (39.36)$$

Such integrals can often not be carried out in closed form, and if the number of nuisance parameters is large, then they can be difficult to compute with standard Monte Carlo methods. *Markov Chain Monte Carlo* (MCMC) techniques are often used for computing integrals of this type (see Sec. 40.5).

#### 39.2.5. Propagation of errors :

Consider a set of  $n$  quantities  $\theta = (\theta_1, \dots, \theta_n)$  and a set of  $m$  functions  $\eta(\theta) = (\eta_1(\theta), \dots, \eta_m(\theta))$ . Suppose we have estimated  $\hat{\theta} = (\hat{\theta}_1, \dots, \hat{\theta}_n)$ , using, say, maximum-likelihood or least-squares, and we also know or have estimated the covariance matrix  $V_{ij} = \text{cov}[\hat{\theta}_i, \hat{\theta}_j]$ . The goal of *error propagation* is to determine the covariance matrix for the functions,  $U_{ij} = \text{cov}[\hat{\eta}_i, \hat{\eta}_j]$ , where  $\hat{\eta} = \eta(\hat{\theta})$ . In particular, the diagonal elements  $U_{ii} = V[\hat{\eta}_i]$  give the variances. The new covariance matrix can be found by expanding the functions  $\eta(\theta)$  about the estimates  $\hat{\theta}$  to first order in a Taylor series. Using this one finds

$$U_{ij} \approx \sum_{k,l} \frac{\partial \eta_i}{\partial \theta_k} \frac{\partial \eta_j}{\partial \theta_l} \Big|_{\hat{\theta}} V_{kl}. \quad (39.37)$$

This can be written in matrix notation as  $U \approx AVA^T$  where the matrix of derivatives  $A$  is

$$A_{ij} = \frac{\partial \eta_i}{\partial \theta_j} \Big|_{\hat{\theta}}, \quad (39.38)$$

and  $A^T$  is its transpose. The approximation is exact if  $\eta(\theta)$  is linear (it holds, for example, in Equation (39.23)). If this is not the case, the approximation can break down if, for example,  $\eta(\theta)$  is significantly nonlinear close to  $\hat{\theta}$  in a region of a size comparable to the standard deviations of  $\hat{\theta}$ .

### 39.3. Statistical tests

In addition to estimating parameters, one often wants to assess the validity of certain statements concerning the data's underlying distribution. Frequentist *hypothesis tests*, described in Sec. 39.3.1, provide a rule for accepting or rejecting hypotheses depending on the outcome of a measurement. In *significance tests*, covered in Sec. 39.3.2, one gives the probability to obtain a level of incompatibility with a certain hypothesis that is greater than or equal to the level observed with the actual data. In the Bayesian approach, the corresponding procedure is based fundamentally on the posterior probabilities of the competing hypotheses. In Sec. 39.3.3 we describe a related construct called the Bayes factor, which can be used to quantify the degree to which the data prefer one or another hypothesis.

#### 39.3.1. Hypothesis tests :

A frequentist *test* of a hypothesis (often called the null hypothesis,  $H_0$ ) is a rule that states for which data values  $\mathbf{x}$  the hypothesis is rejected. A region of  $\mathbf{x}$ -space called the critical region,  $w$ , is specified such that there is no more than a given probability under  $H_0$ ,  $\alpha$ , called the *size* or *significance level* of the test, to find  $\mathbf{x} \in w$ . If the data are discrete, it may not be possible to find a critical region with exact probability content  $\alpha$ , and thus we require  $P(\mathbf{x} \in w|H_0) \leq \alpha$ . If the data are observed in the critical region,  $H_0$  is rejected.

The critical region is not unique. Its choice should take into account the probabilities for the data predicted by some alternative hypothesis (or set of alternatives)  $H_1$ . Rejecting  $H_0$  if it is true is called a *type-I error*, and occurs by construction with probability no greater than  $\alpha$ . Not rejecting  $H_0$  if an alternative  $H_1$  is true is called a *type-II error*, and for a given test this will have a certain probability  $\beta = P(\mathbf{x} \notin w|H_1)$ . The quantity  $1 - \beta$  is called the *power* of the test of  $H_0$  with respect to the alternative  $H_1$ . A strategy for defining the critical region can therefore be to maximize the power with respect to some alternative (or alternatives) given a fixed size  $\alpha$ .

In high-energy physics, the components of  $\mathbf{x}$  might represent the measured properties of candidate events, and the critical region is defined by the cuts that one imposes in order to reject background and thus accept events likely to be of a certain desired type. Here  $H_0$  could represent the background hypothesis and the alternative  $H_1$  could represent the sought after signal. In other cases,  $H_0$  could be the hypothesis that an entire event sample consists of background events only, and the alternative  $H_1$  may represent the hypothesis of a mixture of background and signal.

Often rather than using the full set of quantities  $\mathbf{x}$ , it is convenient to define a scalar function of  $\mathbf{x}$  called a *test statistic*,  $t(\mathbf{x})$ . The critical

region in  $\mathbf{x}$ -space is bounded by a surface of constant  $t(\mathbf{x})$ . Once the function  $t(\mathbf{x})$  is fixed, a given hypothesis for the distribution of  $\mathbf{x}$  will determine a distribution for  $t$ .

To maximize the power of a test of  $H_0$  with respect to the alternative  $H_1$ , the *Neyman–Pearson lemma* states that the critical region  $w$  should be chosen such that for all data values  $\mathbf{x}$  inside  $w$ , the ratio

$$\lambda(\mathbf{x}) = \frac{f(\mathbf{x}|H_1)}{f(\mathbf{x}|H_0)}, \quad (39.39)$$

is greater than a given constant, the value of which is determined by the size of the test  $\alpha$ . Here  $H_0$  and  $H_1$  must be simple hypotheses, i.e., they should not contain undetermined parameters.

The lemma is equivalent to the statement that (39.39) represents the optimal test statistic where the critical region is defined by a single cut on  $\lambda$ . This test will lead to the maximum power for a given probability  $\alpha$  to reject  $H_0$  if  $H_0$  is in fact true. It can be difficult in practice, however, to determine  $\lambda(\mathbf{x})$ , since this requires knowledge of the joint p.d.f.s  $f(\mathbf{x}|H_0)$  and  $f(\mathbf{x}|H_1)$ . Often one does not have explicit formulae for these, but rather Monte Carlo models that allow one to generate instances of  $\mathbf{x}$  (events) that follow the p.d.f.s.

In the case where the likelihood ratio (39.39) cannot be used explicitly, there exist a variety of other multivariate classifiers that effectively separate different types of events. These are based on machine-learning algorithms that use samples of *training data* corresponding to the hypotheses in question, often generated from Monte Carlo models. Methods often used in HEP include *Fisher discriminants* and *neural networks*, *probability density estimation (PDE)* techniques, *kernel-based PDE (KDE or Parzen window)*, *support vector machines*, and *decision trees*. Techniques such as “boosting” and “bagging” can be applied to combine a number of classifiers into a stronger one with greater stability with respect to fluctuations in the training data. Descriptions of these methods can be found in Refs. [11–14], and *Proceedings of the PHYSTAT* conference series [15]. Software for HEP includes the *TMVA* [16], *StatPatternRecognition* [17] and *scikit-learn* [18] packages.

### 39.3.2. Tests of significance (goodness-of-fit) :

Often one wants to quantify the level of agreement between the data and a hypothesis without explicit reference to alternative hypotheses. This can be done by defining a statistic  $t$  that is a function of the data whose value reflects in some way the level of agreement between the data and the hypothesis. The analyst must decide what values of the statistic correspond to better or worse levels of agreement with the hypothesis in question; the choice will in general depend on the relevant alternative hypotheses.

The hypothesis in question,  $H_0$ , will determine the p.d.f.  $f(t|H_0)$  for the statistic. The significance of a discrepancy between the data and what one expects under the assumption of  $H_0$  is quantified by giving the  $p$ -value, defined as the probability to find  $t$  in the region of equal or lesser compatibility with  $H_0$  than the level of compatibility observed with the actual data. For example, if  $t$  is defined such that large values correspond to poor agreement with the hypothesis, then the  $p$ -value would be

$$p = \int_{t_{\text{obs}}}^{\infty} f(t|H_0) dt, \quad (39.40)$$

where  $t_{\text{obs}}$  is the value of the statistic obtained in the actual experiment.

The  $p$ -value should not be confused with the size (significance level) of a test, or the confidence level of a confidence interval (Section 39.4), both of which are pre-specified constants. We may formulate a hypothesis test, however, by defining the critical region to correspond to the data outcomes that give the lowest  $p$ -values, so that finding  $p \leq \alpha$  implies that the data outcome was in the critical region. When constructing a  $p$ -value, one generally chooses the region of data space deemed to have lower compatibility with the model being tested as one having higher compatibility with a given alternative, such that the corresponding test will have a high power with respect to this alternative.

The  $p$ -value is a function of the data, and is therefore itself a random variable. If the hypothesis used to compute the  $p$ -value is true, then for continuous data  $p$  will be uniformly distributed between zero and one. Note that the  $p$ -value is not the probability for the hypothesis; in frequentist statistics, this is not defined.

When searching for a new phenomenon, one tries to reject the hypothesis  $H_0$  that the data are consistent with known (e.g., Standard Model) processes. If the  $p$ -value of  $H_0$  is sufficiently low, then one is willing to accept that some alternative hypothesis is true. Often one converts the  $p$ -value into an equivalent significance  $Z$ , defined so that a  $Z$  standard deviation upward fluctuation of a Gaussian random variable would have an upper tail area equal to  $p$ , i.e.,

$$Z = \Phi^{-1}(1 - p). \quad (39.41)$$

Here  $\Phi$  is the cumulative distribution of the standard Gaussian, and  $\Phi^{-1}$  is its inverse (quantile) function. Often in HEP the level of significance where an effect is said to qualify as a discovery is  $Z = 5$ , i.e., a  $5\sigma$  effect, corresponding to a  $p$ -value of  $2.87 \times 10^{-7}$ . One’s actual degree of belief that a new process is present, however, will depend in general on other factors as well, such as the plausibility of the new signal hypothesis and the degree to which it can describe the data, one’s confidence in the model that led to the observed  $p$ -value, and possible corrections for multiple observations out of which one focuses on the smallest  $p$ -value obtained (the “look-elsewhere effect”, discussed in Section 39.3.2.2).

#### 39.3.2.1. Treatment of nuisance parameters for frequentist tests:

Suppose one wants to test hypothetical values of parameters  $\theta$ , but the model also contains nuisance parameters  $\nu$ . To find a  $p$ -value for  $\theta$  we can construct a test statistic  $q_\theta$  such that larger values constitute increasing incompatibility between the data and the hypothesis. Then for an observed value of the statistic  $q_{\theta, \text{obs}}$ , the  $p$ -value of  $\theta$  is

$$p_\theta(\nu) = \int_{q_{\theta, \text{obs}}}^{\infty} f(q_\theta|\theta, \nu) dq_\theta, \quad (39.42)$$

which depends in general on the nuisance parameters  $\nu$ . In the strict frequentist approach,  $\theta$  is rejected only if the  $p$ -value is less than  $\alpha$  for all possible values of the nuisance parameters.

The difficulty described above is effectively solved if we can define the test statistic  $q_\theta$  in such a way that its distribution  $f(q_\theta|\theta)$  is independent of the nuisance parameters. Although exact independence is only found in special cases, it can be achieved approximately by use of the *profile likelihood ratio*. This is given by the profile likelihood from Eq.(39.18) divided by the value of the likelihood at its maximum, i.e., when evaluated with the ML estimators  $\hat{\theta}$  and  $\hat{\nu}$ :

$$\lambda_p(\theta) = \frac{L(\theta, \hat{\nu}(\theta))}{L(\hat{\theta}, \hat{\nu})}. \quad (39.43)$$

Wilks’ theorem [10] states that, providing certain general conditions are satisfied, the distribution of  $-2 \ln \lambda_p(\theta)$ , under assumption of  $\theta$ , approaches a  $\chi^2$  distribution in the limit where the data sample is very large, independent of the values of the nuisance parameters  $\nu$ . Here the number of degrees of freedom is equal to the number of components of  $\theta$ . More details on use of the profile likelihood are given in Refs. [38–39] and in contributions to the PHYSTAT conferences [15]; explicit formulae for special cases can be found in Ref. 40. Further discussion on how to incorporate systematic uncertainties into  $p$ -values can be found in Ref. 19.

Even with use of the profile likelihood ratio, for a finite data sample the  $p$ -value of hypothesized parameters  $\theta$  will retain in general some dependence on the nuisance parameters  $\nu$ . Ideally one would find the maximum of  $p_\theta(\nu)$  from Eq. (39.42) explicitly, but that is often impractical. An approximate and computationally feasible technique is to use  $p_\theta(\hat{\nu}(\theta))$ , where  $\hat{\nu}(\theta)$  are the profiled values of the nuisance parameters as defined in Section 39.2.2.2. The resulting  $p$ -value is correct if the true values of the nuisance parameters are equal to the profiled values used; otherwise it could be either too high or too low. This is discussed further in Section 39.4.2 on confidence intervals.

One may also treat model uncertainties in a Bayesian manner but then use the resulting model in a frequentist test. Suppose the uncertainty in a set of nuisance parameters  $\nu$  is characterized by a Bayesian prior p.d.f.  $\pi(\nu)$ . This can be used to construct the marginal (also called the prior predictive) model for the data  $\mathbf{x}$  and parameters of interest  $\theta$ ,

$$P_m(\mathbf{x}|\theta) = \int P(\mathbf{x}|\theta, \nu) \pi(\nu) d\nu. \quad (39.44)$$

The marginal model does not represent the probability of data that would be generated if one were really to repeat the experiment, as in that case one would assume that the nuisance parameters do not vary. Rather, the marginal model represents a situation in which every repetition of the experiment is carried out with new values of  $\nu$ , randomly sampled from  $\pi(\nu)$ . It is in effect an average of models each with a given  $\nu$ , where the average is carried out with respect to the prior p.d.f.  $\pi(\nu)$ .

The marginal model for the data  $\mathbf{x}$  can be used to determine the distribution of a test statistic  $Q$ , which can be written

$$P_m(Q|\theta) = \int P(Q|\theta, \nu) \pi(\nu) d\nu. \quad (39.45)$$

In a search for a new signal process, the test statistic can be based on the ratio of likelihoods corresponding to the experiments where signal and background events are both present,  $L_{s+b}$ , to that of background only,  $L_b$ . Often the likelihoods are evaluated with the profiled values of the nuisance parameters, which may give improved performance. It is important to note, however, that it is through use of the marginal model for the distribution of  $Q$  that the uncertainties related to the nuisance parameters are incorporated into the result of the test. Different choices for the test statistic itself only result in variations of the power of the test with respect to different alternatives.

### 39.3.2.2. The look-elsewhere effect:

The “look-elsewhere effect” relates to multiple measurements used to test a single hypothesis. The classic example is when one searches in a distribution for a peak whose position is not predicted in advance. Here the no-peak hypothesis is tested using data in a given range of the distribution. In the frequentist approach the correct  $p$ -value of the no-peak hypothesis is the probability, assuming background only, to find a signal as significant as the one found or more so anywhere in the search region. This can be substantially higher than the probability to find a peak of equal or greater significance in the particular place where it appeared. There is in general some ambiguity as to what constitutes the relevant search region or even the broader set of relevant measurements. Although the desired  $p$ -value is well defined once the search region has been fixed, an exact treatment can require extensive computation.

The “brute-force” solution to this problem by Monte Carlo involves generating data under the background-only hypothesis and for each data set, fitting a peak of unknown position and recording a measure of its significance. To establish a discovery one often requires a  $p$ -value less than  $2.87 \times 10^{-7}$ , corresponding to a  $5\sigma$  or larger effect. Determining this with Monte Carlo thus requires generating and fitting a very large number of experiments, perhaps several times  $10^7$ . In contrast, if the position of the peak is fixed, then the fit to the distribution is much easier, and furthermore one can in many cases use formulae valid for sufficiently large samples that bypass completely the need for Monte Carlo (see, e.g., [40]). However, this fixed-position or “local”  $p$ -value would not be correct in general, as it assumes the position of the peak was known in advance.

A method that allows one to modify the local  $p$ -value computed under assumption of a fixed position to obtain an approximation to the correct “global” value using a relatively simple calculation is described in Ref. 20. Suppose a test statistic  $q_0$ , defined so that larger values indicate increasing disagreement with the data, is observed to have a value  $u$ . Furthermore suppose the model contains a nuisance parameter  $\theta$  (such as the peak position) which is only defined under the signal model (there is no peak in the background-only model). An approximation for the global  $p$ -value is found to be

$$p_{\text{global}} \approx p_{\text{local}} + \langle N_u \rangle, \quad (39.46)$$

where  $\langle N_u \rangle$  is the mean number of “upcrossings” of the statistic  $q_0$  above the level  $u$  in the range of the nuisance parameter considered (e.g., the mass range).

The value of  $\langle N_u \rangle$  can be estimated from the number of upcrossings  $\langle N_{u_0} \rangle$  above some much lower value,  $u_0$ , by using a relation due to Davis [21],

$$\langle N_u \rangle \approx \langle N_{u_0} \rangle e^{-(u-u_0)/2}. \quad (39.47)$$

By choosing  $u_0$  sufficiently low, the value of  $\langle N_u \rangle$  can be estimated by simulating only a very small number of experiments, or even from the observed data, rather than the  $10^7$  needed if one is dealing with a  $5\sigma$  effect.

### 39.3.2.3. Goodness-of-fit with the method of Least Squares:

When estimating parameters using the method of least squares, one obtains the minimum value of the quantity  $\chi^2$  (39.19). This statistic can be used to test the *goodness-of-fit*, i.e., the test provides a measure of the significance of a discrepancy between the data and the hypothesized functional form used in the fit. It may also happen that no parameters are estimated from the data, but that one simply wants to compare a histogram, e.g., a vector of Poisson distributed numbers  $\mathbf{n} = (n_1, \dots, n_N)$ , with a hypothesis for their expectation values  $\mu_i = E[n_i]$ . As the distribution is Poisson with variances  $\sigma_i^2 = \mu_i$ , the  $\chi^2$  (39.19) becomes *Pearson's  $\chi^2$  statistic*,

$$\chi^2 = \sum_{i=1}^N \frac{(n_i - \mu_i)^2}{\mu_i}. \quad (39.48)$$

If the hypothesis  $\mu = (\mu_1, \dots, \mu_N)$  is correct, and if the expected values  $\mu_i$  in (39.48) are sufficiently large (or equivalently, if the measurements  $n_i$  can be treated as following a Gaussian distribution), then the  $\chi^2$  statistic will follow the  $\chi^2$  p.d.f. with the number of degrees of freedom equal to the number of measurements  $N$  minus the number of fitted parameters.

Alternatively, one may fit parameters and evaluate goodness-of-fit by minimizing  $-2 \ln \lambda$  from Eq. (39.16). One finds that the distribution of this statistic approaches the asymptotic limit faster than does Pearson's  $\chi^2$ , and thus computing the  $p$ -value with the  $\chi^2$  p.d.f. will in general be better justified (see Ref. 9 and references therein).

Assuming the goodness-of-fit statistic follows a  $\chi^2$  p.d.f., the  $p$ -value for the hypothesis is then

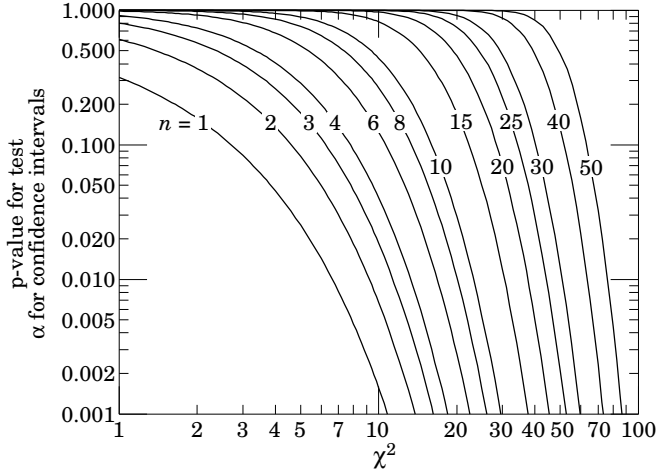
$$p = \int_{\chi^2}^{\infty} f(z; n_d) dz, \quad (39.49)$$

where  $f(z; n_d)$  is the  $\chi^2$  p.d.f. and  $n_d$  is the appropriate number of degrees of freedom. Values are shown in Fig. 39.1 or obtained from the ROOT function `TMath::Prob`. If the conditions for using the  $\chi^2$  p.d.f. do not hold, the statistic can still be defined as before, but its p.d.f. must be determined by other means in order to obtain the  $p$ -value, e.g., using a Monte Carlo calculation.

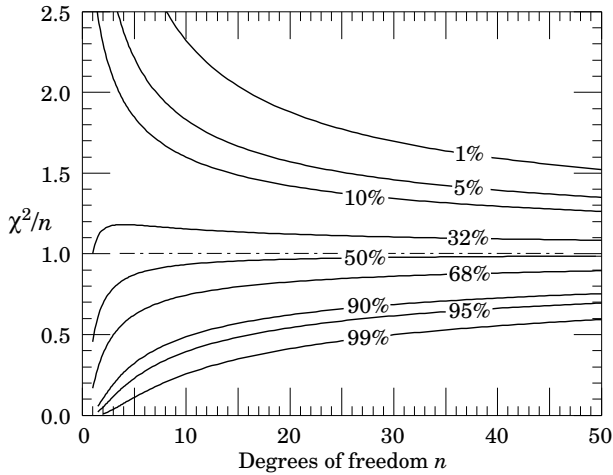
Since the mean of the  $\chi^2$  distribution is equal to  $n_d$ , one expects in a “reasonable” experiment to obtain  $\chi^2 \approx n_d$ . Hence the quantity  $\chi^2/n_d$  is sometimes reported. Since the p.d.f. of  $\chi^2/n_d$  depends on  $n_d$ , however, one must report  $n_d$  as well if one wishes to determine the  $p$ -value. The  $p$ -values obtained for different values of  $\chi^2/n_d$  are shown in Fig. 39.2.

If one finds a  $\chi^2$  value much greater than  $n_d$ , and a correspondingly small  $p$ -value, one may be tempted to expect a high degree of uncertainty for any fitted parameters. Poor goodness-of-fit, however, does not mean that one will have large statistical errors for parameter estimates. If, for example, the error bars (or covariance matrix) used in constructing the  $\chi^2$  are underestimated, then this will lead to underestimated statistical errors for the fitted parameters. The standard deviations of estimators that one finds from, say, Eq. (39.13) reflect how widely the estimates would be distributed if one were to repeat the measurement many times, assuming that the hypothesis and measurement errors used in the  $\chi^2$  are also correct. They do not include the systematic error which may result from an incorrect hypothesis or incorrectly estimated measurement errors in the  $\chi^2$ .





**Figure 39.1:** One minus the  $\chi^2$  cumulative distribution,  $1 - F(\chi^2; n)$ , for  $n$  degrees of freedom. This gives the  $p$ -value for the  $\chi^2$  goodness-of-fit test as well as one minus the coverage probability for confidence regions (see Sec. 39.4.2.2).



**Figure 39.2:** The ‘reduced’  $\chi^2$ , equal to  $\chi^2/n$ , for  $n$  degrees of freedom. The curves show as a function of  $n$  the  $\chi^2/n$  that corresponds to a given  $p$ -value.

### 39.3.3. Bayes factors :

In Bayesian statistics, all of one’s knowledge about a model is contained in its posterior probability, which one obtains using Bayes’ theorem (Eq. (39.30)). Thus one could reject a hypothesis  $H$  if its posterior probability  $P(H|\mathbf{x})$  is sufficiently small. The difficulty here is that  $P(H|\mathbf{x})$  is proportional to the prior probability  $P(H)$ , and there will not be a consensus about the prior probabilities for the existence of new phenomena. Nevertheless one can construct a quantity called the Bayes factor (described below), which can be used to quantify the degree to which the data prefer one hypothesis over another, and is independent of their prior probabilities.

Consider two models (hypotheses),  $H_i$  and  $H_j$ , described by vectors of parameters  $\theta_i$  and  $\theta_j$ , respectively. Some of the components will be common to both models and others may be distinct. The full prior probability for each model can be written in the form

$$\pi(H_i, \theta_i) = P(H_i)\pi(\theta_i|H_i). \quad (39.50)$$

Here  $P(H_i)$  is the overall prior probability for  $H_i$ , and  $\pi(\theta_i|H_i)$  is the normalized p.d.f. of its parameters. For each model, the posterior probability is found using Bayes’ theorem,

$$P(H_i|\mathbf{x}) = \frac{\int P(\mathbf{x}|\theta_i, H_i)P(H_i)\pi(\theta_i|H_i) d\theta_i}{P(\mathbf{x})}, \quad (39.51)$$

where the integration is carried out over the internal parameters  $\theta_i$  of the model. The ratio of posterior probabilities for the models is therefore

$$\frac{P(H_i|\mathbf{x})}{P(H_j|\mathbf{x})} = \frac{\int P(\mathbf{x}|\theta_i, H_i)\pi(\theta_i|H_i) d\theta_i}{\int P(\mathbf{x}|\theta_j, H_j)\pi(\theta_j|H_j) d\theta_j} \frac{P(H_i)}{P(H_j)}. \quad (39.52)$$

The *Bayes factor* is defined as

$$B_{ij} = \frac{\int P(\mathbf{x}|\theta_i, H_i)\pi(\theta_i|H_i) d\theta_i}{\int P(\mathbf{x}|\theta_j, H_j)\pi(\theta_j|H_j) d\theta_j}. \quad (39.53)$$

This gives what the ratio of posterior probabilities for models  $i$  and  $j$  would be if the overall prior probabilities for the two models were equal. If the models have no nuisance parameters, *i.e.*, no internal parameters described by priors, then the Bayes factor is simply the likelihood ratio. The Bayes factor therefore shows by how much the probability ratio of model  $i$  to model  $j$  changes in the light of the data, and thus can be viewed as a numerical measure of evidence supplied by the data in favour of one hypothesis over the other.

Although the Bayes factor is by construction independent of the overall prior probabilities  $P(H_i)$  and  $P(H_j)$ , it does require priors for all internal parameters of a model, *i.e.*, one needs the functions  $\pi(\theta_i|H_i)$  and  $\pi(\theta_j|H_j)$ . In a Bayesian analysis where one is only interested in the posterior p.d.f. of a parameter, it may be acceptable to take an unnormalizable function for the prior (an improper prior) as long as the product of likelihood and prior can be normalized. But improper priors are only defined up to an arbitrary multiplicative constant, and so the Bayes factor would depend on this constant. Furthermore, although the range of a constant normalized prior is unimportant for parameter determination (provided it is wider than the likelihood), this is not so for the Bayes factor when such a prior is used for only one of the hypotheses. So to compute a Bayes factor, all internal parameters must be described by normalized priors that represent meaningful probabilities over the entire range where they are defined.

An exception to this rule may be considered when the identical parameter appears in the models for both numerator and denominator of the Bayes factor. In this case one can argue that the arbitrary constants would cancel. One must exercise some caution, however, as parameters with the same name and physical meaning may still play different roles in the two models.

Both integrals in Equation (39.53) are of the form

$$m = \int P(\mathbf{x}|\theta)\pi(\theta) d\theta, \quad (39.54)$$

which is the marginal likelihood seen previously in Eq. (39.44) (in some fields this quantity is called the *evidence*). A review of Bayes factors can be found in Ref. 32. Computing marginal likelihoods can be difficult; in many cases it can be done with the nested sampling algorithm [33] as implemented, *e.g.*, in the program `MultiNest` [34].

## 39.4. Intervals and limits

When the goal of an experiment is to determine a parameter  $\theta$ , the result is usually expressed by quoting, in addition to the point estimate, some sort of interval which reflects the statistical precision of the measurement. In the simplest case, this can be given by the parameter’s estimated value  $\hat{\theta}$  plus or minus an estimate of the standard deviation of  $\hat{\theta}$ ,  $\hat{\sigma}_{\hat{\theta}}$ . If, however, the p.d.f. of the estimator is not Gaussian or if there are physical boundaries on the possible values of the parameter, then one usually quotes instead an interval according to one of the procedures described below.

In reporting an interval or limit, the experimenter may wish to

- communicate as objectively as possible the result of the experiment;
- provide an interval that is constructed to cover the true value of the parameter with a specified probability;
- provide the information needed by the consumer of the result to draw conclusions about the parameter or to make a particular decision;
- draw conclusions about the parameter that incorporate stated prior beliefs.

With a sufficiently large data sample, the point estimate and standard deviation (or for the multiparameter case, the parameter estimates and covariance matrix) satisfy essentially all of these goals. For finite data samples, no single method for quoting an interval will achieve all of them.

In addition to the goals listed above, the choice of method may be influenced by practical considerations such as ease of producing an interval from the results of several measurements. Of course the experimenter is not restricted to quoting a single interval or limit; one may choose, for example, first to communicate the result with a confidence interval having certain frequentist properties, and then in addition to draw conclusions about a parameter using a judiciously chosen subjective Bayesian prior. It is recommended, however, that there be a clear separation between these two aspects of reporting a result. In the remainder of this section, we assess the extent to which various types of intervals achieve the goals stated here.

#### 39.4.1. Bayesian intervals :

As described in Sec. 39.2.4, a Bayesian posterior probability may be used to determine regions that will have a given probability of containing the true value of a parameter. In the single parameter case, for example, an interval (called a Bayesian or credible interval)  $[\theta_{\text{lo}}, \theta_{\text{up}}]$  can be determined which contains a given fraction  $1 - \alpha$  of the posterior probability, *i.e.*,

$$1 - \alpha = \int_{\theta_{\text{lo}}}^{\theta_{\text{up}}} p(\theta|\mathbf{x}) d\theta. \quad (39.55)$$

Sometimes an upper or lower limit is desired, *i.e.*,  $\theta_{\text{lo}}$  or  $\theta_{\text{up}}$  can be set to a physical boundary or to plus or minus infinity. In other cases, one might be interested in the set of  $\theta$  values for which  $p(\theta|\mathbf{x})$  is higher than for any  $\theta$  not belonging to the set, which may constitute a single interval or a set of disjoint regions; these are called highest posterior density (HPD) intervals. Note that HPD intervals are not invariant under a nonlinear transformation of the parameter.

If a parameter is constrained to be non-negative, then the prior p.d.f. can simply be set to zero for negative values. An important example is the case of a Poisson variable  $n$ , which counts signal events with unknown mean  $s$ , as well as background with mean  $b$ , assumed known. For the signal mean  $s$ , one often uses the prior

$$\pi(s) = \begin{cases} 0 & s < 0 \\ 1 & s \geq 0 \end{cases}. \quad (39.56)$$

This prior is regarded as providing an interval whose frequentist properties can be studied, rather than as representing a degree of belief. For example, to obtain an upper limit on  $s$ , one may proceed as follows. The likelihood for  $s$  is given by the Poisson distribution for  $n$  with mean  $s + b$ ,

$$P(n|s) = \frac{(s+b)^n}{n!} e^{-(s+b)}, \quad (39.57)$$

along with the prior (39.56) in (39.30) gives the posterior density for  $s$ . An upper limit  $s_{\text{up}}$  at confidence level (or here, rather, credibility level)  $1 - \alpha$  can be obtained by requiring

$$1 - \alpha = \int_{-\infty}^{s_{\text{up}}} p(s|n) ds = \frac{\int_{-\infty}^{s_{\text{up}}} P(n|s) \pi(s) ds}{\int_{-\infty}^{\infty} P(n|s) \pi(s) ds}, \quad (39.58)$$

where the lower limit of integration is effectively zero because of the cut-off in  $\pi(s)$ . By relating the integrals in Eq. (39.58) to incomplete gamma functions, the solution for the upper limit is found to be

$$s_{\text{up}} = \frac{1}{2} F_{\chi^2}^{-1}[p, 2(n+1)] - b, \quad (39.59)$$

where  $F_{\chi^2}^{-1}$  is the quantile of the  $\chi^2$  distribution (inverse of the cumulative distribution). Here the quantity  $p$  is

$$p = 1 - \alpha \left( 1 - F_{\chi^2}[2b, 2(n+1)] \right), \quad (39.60)$$

where  $F_{\chi^2}$  is the cumulative  $\chi^2$  distribution. For both  $F_{\chi^2}$  and  $F_{\chi^2}^{-1}$  above, the argument  $2(n+1)$  gives the number of degrees of freedom. For the special case of  $b = 0$ , the limit reduces to

$$s_{\text{up}} = \frac{1}{2} F_{\chi^2}^{-1}(1 - \alpha; 2(n+1)). \quad (39.61)$$

It happens that for the case of  $b = 0$ , the upper limit from Eq. (39.61) coincides numerically with the frequentist upper limit discussed in Section 39.4.2.3. Values for  $1 - \alpha = 0.9$  and  $0.95$  are given by the values  $\mu_{\text{up}}$  in Table 39.3. The frequentist properties of confidence intervals for the Poisson mean found in this way are discussed in Refs. [2] and [23].

As in any Bayesian analysis, it is important to show how the result changes under assumption of different prior probabilities. For example, one could consider the Jeffreys prior as described in Sec. 39.2.4. For this problem one finds the Jeffreys prior  $\pi(s) \propto 1/\sqrt{s+b}$  for  $s \geq 0$  and zero otherwise. As with the constant prior, one would not regard this as representing one's prior beliefs about  $s$ , both because it is improper and also as it depends on  $b$ . Rather it is used with Bayes' theorem to produce an interval whose frequentist properties can be studied.

If the model contains nuisance parameters then these are eliminated by marginalizing, as in Eq. (39.36), to obtain the p.d.f. for the parameters of interest. For example, if the parameter  $b$  in the Poisson counting problem above were to be characterized by a prior p.d.f.  $\pi(b)$ , then one would first use Bayes' theorem to find  $p(s, b|n)$ . This is then marginalized to find  $p(s|n) = \int p(s, b|n) \pi(b) db$ , from which one may determine an interval for  $s$ . One may not be certain whether to extend a model by including more nuisance parameters. In this case, a Bayes factor may be used to determine to what extent the data prefer a model with additional parameters, as described in Section 39.3.3.

#### 39.4.2. Frequentist confidence intervals :

The unqualified phrase "confidence intervals" refers to frequentist intervals obtained with a procedure due to Neyman [31], described below. These are intervals (or in the multiparameter case, regions) constructed so as to include the true value of the parameter with a probability greater than or equal to a specified level, called the *coverage probability*. It is important to note that in the frequentist approach, such coverage is not meaningful for a fixed interval. A confidence interval, however, depends on the data and thus would fluctuate if one were to repeat the experiment many times. The coverage probability refers to the fraction of intervals in such a set that contain the true parameter value. In this section we discuss several techniques for producing intervals that have, at least approximately, this property.

##### 39.4.2.1. The Neyman construction for confidence intervals:

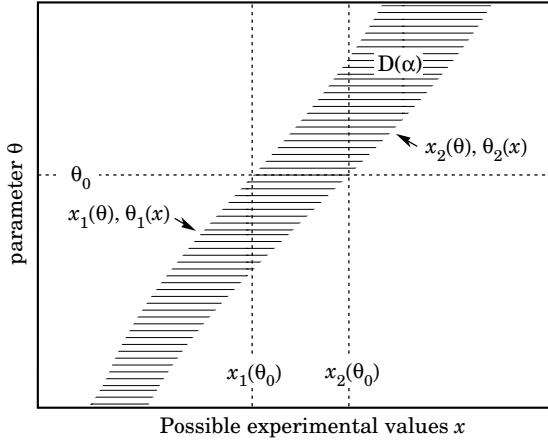
Consider a p.d.f.  $f(x; \theta)$  where  $x$  represents the outcome of the experiment and  $\theta$  is the unknown parameter for which we want to construct a confidence interval. The variable  $x$  could (and often does) represent an estimator for  $\theta$ . Using  $f(x; \theta)$ , we can find for a pre-specified probability  $1 - \alpha$ , and for every value of  $\theta$ , a set of values  $x_1(\theta, \alpha)$  and  $x_2(\theta, \alpha)$  such that

$$P(x_1 < x < x_2; \theta) = \int_{x_1}^{x_2} f(x; \theta) dx \geq 1 - \alpha. \quad (39.62)$$

If  $x$  is discrete, the integral is replaced by the corresponding sum. In that case there may not exist a range of  $x$  values whose summed probability is exactly equal to a given value of  $1 - \alpha$ , and one requires by convention  $P(x_1 < x < x_2; \theta) \geq 1 - \alpha$ .

This is illustrated for continuous  $x$  in Fig. 39.3: a horizontal line segment  $[x_1(\theta, \alpha), x_2(\theta, \alpha)]$  is drawn for representative values of  $\theta$ . The union of such intervals for all values of  $\theta$ , designated in the figure as  $D(\alpha)$ , is known as the *confidence belt*. Typically the curves  $x_1(\theta, \alpha)$  and  $x_2(\theta, \alpha)$  are monotonic functions of  $\theta$ , which we assume for this discussion.

Upon performing an experiment to measure  $x$  and obtaining a value  $x_0$ , one draws a vertical line through  $x_0$ . The confidence interval for  $\theta$  is the set of all values of  $\theta$  for which the corresponding line segment



**Figure 39.3:** Construction of the confidence belt (see text).

$[x_1(\theta, \alpha), x_2(\theta, \alpha)]$  is intercepted by this vertical line. Such confidence intervals are said to have a *confidence level* (CL) equal to  $1 - \alpha$ .

Now suppose that the true value of  $\theta$  is  $\theta_0$ , indicated in the figure. We see from the figure that  $\theta_0$  lies between  $\theta_1(x)$  and  $\theta_2(x)$  if and only if  $x$  lies between  $x_1(\theta_0)$  and  $x_2(\theta_0)$ . The two events thus have the same probability, and since this is true for any value  $\theta_0$ , we can drop the subscript 0 and obtain

$$1 - \alpha = P(x_1(\theta) < x < x_2(\theta)) = P(\theta_2(x) < \theta < \theta_1(x)). \quad (39.63)$$

In this probability statement,  $\theta_1(x)$  and  $\theta_2(x)$ , *i.e.*, the endpoints of the interval, are the random variables and  $\theta$  is an unknown constant. If the experiment were to be repeated a large number of times, the interval  $[\theta_1, \theta_2]$  would vary, covering the fixed value  $\theta$  in a fraction  $1 - \alpha$  of the experiments.

The condition of coverage in Eq. (39.62) does not determine  $x_1$  and  $x_2$  uniquely, and additional criteria are needed. One possibility is to choose *central intervals* such that the probabilities to find  $x$  below  $x_1$  and above  $x_2$  are each  $\alpha/2$ . In other cases, one may want to report only an upper or lower limit, in which case one of  $P(x \leq x_1)$  or  $P(x \geq x_2)$  can be set to  $\alpha$  and the other to zero. Another principle based on *likelihood ratio ordering* for determining which values of  $x$  should be included in the confidence belt is discussed below.

When the observed random variable  $x$  is continuous, the coverage probability obtained with the Neyman construction is  $1 - \alpha$ , regardless of the true value of the parameter. Because of the requirement  $P(x_1 < x < x_2) \geq 1 - \alpha$  when  $x$  is discrete, one obtains in that case confidence intervals that include the true parameter with a probability greater than or equal to  $1 - \alpha$ .

An equivalent method of constructing confidence intervals is to consider a test (see Sec. 39.3) of the hypothesis that the parameter's true value is  $\theta$  (assume one constructs a test for all physical values of  $\theta$ ). One then excludes all values of  $\theta$  where the hypothesis would be rejected in a test of size  $\alpha$  or less. The remaining values constitute the confidence interval at confidence level  $1 - \alpha$ . If the critical region of the test is characterized by having a  $p$ -value  $p_\theta \leq \alpha$ , then the endpoints of the confidence interval are found in practice by solving  $p_\theta = \alpha$  for  $\theta$ .

In the procedure outlined above, one is still free to choose the test to be used; this corresponds to the freedom in the Neyman construction as to which values of the data are included in the confidence belt. One possibility is to use a test statistic based on the *likelihood ratio*,

$$\lambda(\theta) = \frac{f(x; \theta)}{f(x; \hat{\theta})}, \quad (39.64)$$

where  $\hat{\theta}$  is the value of the parameter which, out of all allowed values, maximizes  $f(x; \theta)$ . This results in the intervals described in Ref. 35 by Feldman and Cousins. The same intervals can be obtained from the Neyman construction described above by including in the confidence belt those values of  $x$  which give the greatest values of  $\lambda(\theta)$ .

If the model contains nuisance parameters  $\nu$ , then these can be incorporated into the test (or the  $p$ -values) used to determine the limit by profiling as discussed in Section 39.3.2.1. As mentioned there, the strict frequentist approach is to regard the parameter of interest  $\theta$  as excluded only if it is rejected for all possible values of  $\nu$ . The resulting interval for  $\theta$  will then cover the true value with a probability greater than or equal to the nominal confidence level for all points in  $\nu$ -space.

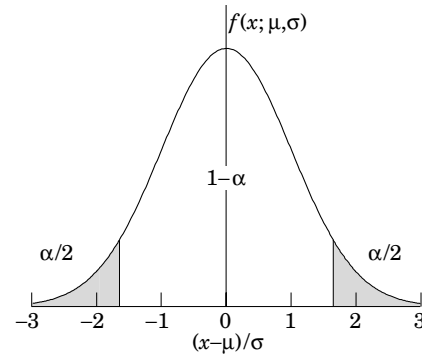
If the  $p$ -value is based on the profiled values of the nuisance parameters, *i.e.*, with  $\nu = \hat{\nu}(\theta)$  used in Eq. (39.42), then the resulting interval for the parameter of interest will have the correct coverage if the true values of  $\nu$  are equal to the profiled values. Otherwise the coverage probability may be too high or too low. This procedure has been called *profile construction* in HEP [22] (see also [19]).

#### 39.4.2.2. Gaussian distributed measurements:

An important example of constructing a confidence interval is when the data consists of a single random variable  $x$  that follows a Gaussian distribution; this is often the case when  $x$  represents an estimator for a parameter and one has a sufficiently large data sample. If there is more than one parameter being estimated, the multivariate Gaussian is used. For the univariate case with known  $\sigma$ , the probability that the measured value  $x$  will fall within  $\pm\delta$  of the true value  $\mu$  is

$$1 - \alpha = \frac{1}{\sqrt{2\pi}\sigma} \int_{\mu-\delta}^{\mu+\delta} e^{-(x-\mu)^2/2\sigma^2} dx = \text{erf}\left(\frac{\delta}{\sqrt{2}\sigma}\right) = 2\Phi\left(\frac{\delta}{\sigma}\right) - 1, \quad (39.65)$$

where erf is the Gaussian error function, which is rewritten in the final equality using  $\Phi$ , the Gaussian cumulative distribution. Fig. 39.4 shows a  $\delta = 1.64\sigma$  confidence interval unshaded. The choice  $\delta = \sigma$  gives an interval called the *standard error* which has  $1 - \alpha = 68.27\%$  if  $\sigma$  is known. Values of  $\alpha$  for other frequently used choices of  $\delta$  are given in Table 39.1.



**Figure 39.4:** Illustration of a symmetric 90% confidence interval (unshaded) for a measurement of a single quantity with Gaussian errors. Integrated probabilities, defined by  $\alpha = 0.1$ , are as shown.

**Table 39.1:** Area of the tails  $\alpha$  outside  $\pm\delta$  from the mean of a Gaussian distribution.

$\alpha$	$\delta$	$\alpha$	$\delta$
0.3173	$1\sigma$	0.2	$1.28\sigma$
$4.55 \times 10^{-2}$	$2\sigma$	0.1	$1.64\sigma$
$2.7 \times 10^{-3}$	$3\sigma$	0.05	$1.96\sigma$
$6.3 \times 10^{-5}$	$4\sigma$	0.01	$2.58\sigma$
$5.7 \times 10^{-7}$	$5\sigma$	0.001	$3.29\sigma$
$2.0 \times 10^{-9}$	$6\sigma$	$10^{-4}$	$3.89\sigma$

We can set a one-sided (upper or lower) limit by excluding above  $x + \delta$  (or below  $x - \delta$ ). The values of  $\alpha$  for such limits are half the values in Table 39.1.

The relation (39.65) can be re-expressed using the cumulative distribution function for the  $\chi^2$  distribution as

$$\alpha = 1 - F(\chi^2; n), \quad (39.66)$$

for  $\chi^2 = (\delta/\sigma)^2$  and  $n = 1$  degree of freedom. This can be seen as the  $n = 1$  curve in Fig. 39.1 or obtained by using the ROOT function `TMath::Prob`.

For multivariate measurements of, say,  $n$  parameter estimates  $\hat{\theta} = (\hat{\theta}_1, \dots, \hat{\theta}_n)$ , one requires the full covariance matrix  $V_{ij} = \text{cov}[\hat{\theta}_i, \hat{\theta}_j]$ , which can be estimated as described in Sections 39.2.2 and 39.2.3. Under fairly general conditions with the methods of maximum-likelihood or least-squares in the large sample limit, the estimators will be distributed according to a multivariate Gaussian centered about the true (unknown) values  $\theta$ , and furthermore, the likelihood function itself takes on a Gaussian shape.

The standard error ellipse for the pair  $(\hat{\theta}_i, \hat{\theta}_j)$  is shown in Fig. 39.5, corresponding to a contour  $\chi^2 = \chi_{\min}^2 + 1$  or  $\ln L = \ln L_{\max} - 1/2$ .

The ellipse is centered about the estimated values  $\hat{\theta}$ , and the tangents to the ellipse give the standard deviations of the estimators,  $\sigma_i$  and  $\sigma_j$ . The angle of the major axis of the ellipse is given by

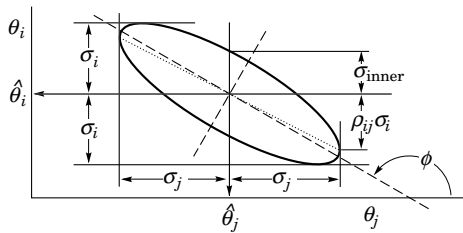
$$\tan 2\phi = \frac{2\rho_{ij}\sigma_i\sigma_j}{\sigma_j^2 - \sigma_i^2}, \quad (39.67)$$

where  $\rho_{ij} = \text{cov}[\hat{\theta}_i, \hat{\theta}_j]/\sigma_i\sigma_j$  is the correlation coefficient.

The correlation coefficient can be visualized as the fraction of the distance  $\sigma_i$  from the ellipse's horizontal center-line at which the ellipse becomes tangent to vertical, *i.e.*, at the distance  $\rho_{ij}\sigma_i$  below the center-line as shown. As  $\rho_{ij}$  goes to  $+1$  or  $-1$ , the ellipse thins to a diagonal line.

It could happen that one of the parameters, say,  $\theta_j$ , is known from previous measurements to a precision much better than  $\sigma_j$ , so that the current measurement contributes almost nothing to the knowledge of  $\theta_j$ . However, the current measurement of  $\theta_i$  and its dependence on  $\theta_j$  may still be important. In this case, instead of quoting both parameter estimates and their correlation, one sometimes reports the value of  $\theta_i$ , which minimizes  $\chi^2$  at a fixed value of  $\theta_j$ , such as the PDG best value. This  $\theta_i$  value lies along the dotted line between the points where the ellipse becomes tangent to vertical, and has statistical error  $\sigma_{\text{inner}}$  as shown on the figure, where  $\sigma_{\text{inner}} = (1 - \rho_{ij}^2)^{1/2}\sigma_i$ .

Instead of the correlation  $\rho_{ij}$ , one reports the dependency  $d\hat{\theta}_i/d\theta_j$ , which is the slope of the dotted line. This slope is related to the correlation coefficient by  $d\hat{\theta}_i/d\theta_j = \rho_{ij} \times \frac{\sigma_i}{\sigma_j}$ .



**Figure 39.5:** Standard error ellipse for the estimators  $\hat{\theta}_i$  and  $\hat{\theta}_j$ . In the case shown the correlation is negative.

As in the single-variable case, because of the symmetry of the Gaussian function between  $\theta$  and  $\hat{\theta}$ , one finds that contours of constant  $\ln L$  or  $\chi^2$  cover the true values with a certain, fixed probability. That is, the confidence region is determined by

$$\ln L(\theta) \geq \ln L_{\max} - \Delta \ln L, \quad (39.68)$$

or where a  $\chi^2$  has been defined for use with the method of least-squares,

$$\chi^2(\theta) \leq \chi_{\min}^2 + \Delta \chi^2. \quad (39.69)$$

Values of  $\Delta \chi^2$  or  $2\Delta \ln L$  are given in Table 39.2 for several values of the coverage probability  $1 - \alpha$  and number of fitted parameters  $m$ . For Gaussian distributed data, these are related by  $\Delta \chi^2 = 2\Delta \ln L = F_{\chi_m^2}^{-1}(1 - \alpha)$ , where  $F_{\chi_m^2}^{-1}$  is the chi-square quantile (inverse of the cumulative distribution) for  $m$  degrees of freedom.

**Table 39.2:** Values of  $\Delta \chi^2$  or  $2\Delta \ln L$  corresponding to a coverage probability  $1 - \alpha$  in the large data sample limit, for joint estimation of  $m$  parameters.

$(1 - \alpha) (\%)$	$m = 1$	$m = 2$	$m = 3$
68.27	1.00	2.30	3.53
90.	2.71	4.61	6.25
95.	3.84	5.99	7.82
95.45	4.00	6.18	8.03
99.	6.63	9.21	11.34
99.73	9.00	11.83	14.16

For non-Gaussian data samples, the probability for the regions determined by Equations (39.68) or (39.69) to cover the true value of  $\theta$  becomes independent of  $\theta$  only in the large-sample limit. So for a finite data sample these are not exact confidence regions according to our previous definition. Nevertheless, they can still have a coverage probability only weakly dependent on the true parameter, and approximately as given in Table 39.2. In any case, the coverage probability of the intervals or regions obtained according to this procedure can in principle be determined as a function of the true parameter(s), for example, using a Monte Carlo calculation.

One of the practical advantages of intervals that can be constructed from the log-likelihood function or  $\chi^2$  is that it is relatively simple to produce the interval for the combination of several experiments. If  $N$  independent measurements result in log-likelihood functions  $\ln L_i(\theta)$ , then the combined log-likelihood function is simply the sum,

$$\ln L(\theta) = \sum_{i=1}^N \ln L_i(\theta). \quad (39.70)$$

This can then be used to determine an approximate confidence interval or region with Eq. (39.68), just as with a single experiment.

#### 39.4.2.3. Poisson or binomial data:

Another important class of measurements consists of counting a certain number of events,  $n$ . In this section, we will assume these are all events of the desired type, *i.e.*, there is no background. If  $n$  represents the number of events produced in a reaction with cross section  $\sigma$ , say, in a fixed integrated luminosity  $\mathcal{L}$ , then it follows a Poisson distribution with mean  $\mu = \sigma \mathcal{L}$ . If, on the other hand, one has selected a larger sample of  $N$  events and found  $n$  of them to have a particular property, then  $n$  follows a binomial distribution where the parameter  $p$  gives the probability for the event to possess the property in question. This is appropriate, *e.g.*, for estimates of branching ratios or selection efficiencies based on a given total number of events.

For the case of Poisson distributed  $n$ , limits on the mean value  $\mu$  can be found from the Neyman procedure as discussed in Section 39.4.2.1 with  $n$  used directly as the statistic  $x$ . The upper and lower limits are found to be

$$\mu_{\text{lo}} = \frac{1}{2} F_{\chi^2}^{-1}(\alpha_{\text{lo}}; 2n), \quad (39.71a)$$

$$\mu_{\text{up}} = \frac{1}{2} F_{\chi^2}^{-1}(1 - \alpha_{\text{up}}; 2(n+1)), \quad (39.71b)$$

where confidence levels of  $1 - \alpha_{\text{lo}}$  and  $1 - \alpha_{\text{up}}$ , refer separately to the corresponding intervals  $\mu \geq \mu_{\text{lo}}$  and  $\mu \leq \mu_{\text{up}}$ , and  $F_{\chi^2}^{-1}$  is the quantile of the  $\chi^2$  distribution (inverse of the cumulative distribution). The quantiles  $F_{\chi^2}^{-1}$  can be obtained from standard tables or from the ROOT routine `TMath::ChisquareQuantile`. For central confidence intervals at confidence level  $1 - \alpha$ , set  $\alpha_{\text{lo}} = \alpha_{\text{up}} = \alpha/2$ .

It happens that the upper limit from Eq. (39.71b) coincides numerically with the Bayesian upper limit for a Poisson parameter, using a uniform prior p.d.f. for  $\mu$ . Values for confidence levels of 90% and 95% are shown in Table 39.3. For the case of binomially distributed  $n$  successes out of  $N$  trials with probability of success  $p$ ,

**Table 39.3:** Lower and upper (one-sided) limits for the mean  $\mu$  of a Poisson variable given  $n$  observed events in the absence of background, for confidence levels of 90% and 95%.

$n$	$1 - \alpha = 90\%$		$1 - \alpha = 95\%$	
	$\mu_{\text{lo}}$	$\mu_{\text{up}}$	$\mu_{\text{lo}}$	$\mu_{\text{up}}$
0	–	2.30	–	3.00
1	0.105	3.89	0.051	4.74
2	0.532	5.32	0.355	6.30
3	1.10	6.68	0.818	7.75
4	1.74	7.99	1.37	9.15
5	2.43	9.27	1.97	10.51
6	3.15	10.53	2.61	11.84
7	3.89	11.77	3.29	13.15
8	4.66	12.99	3.98	14.43
9	5.43	14.21	4.70	15.71
10	6.22	15.41	5.43	16.96

the upper and lower limits on  $p$  are found to be

$$p_{\text{lo}} = \frac{nF_F^{-1}[\alpha_{\text{lo}}; 2n, 2(N - n + 1)]}{N - n + 1 + nF_F^{-1}[\alpha_{\text{lo}}; 2n, 2(N - n + 1)]}, \quad (39.72a)$$

$$p_{\text{up}} = \frac{(n + 1)F_F^{-1}[1 - \alpha_{\text{up}}; 2(n + 1), 2(N - n)]}{(N - n) + (n + 1)F_F^{-1}[1 - \alpha_{\text{up}}; 2(n + 1), 2(N - n)]}. \quad (39.72b)$$

Here  $F_F^{-1}$  is the quantile of the  $F$  distribution (also called the Fisher–Snedecor distribution; see Ref. 4).

#### 39.4.2.4. Parameter exclusion in cases of low sensitivity:

An important example of a statistical test arises in the search for a new signal process. Suppose the parameter  $\mu$  is defined such that it is proportional to the signal cross section. A statistical test may be carried out for hypothesized values of  $\mu$ , which may be done by computing a  $p$ -value,  $p_\mu$ , for all  $\mu$ . Those values not rejected in a test of size  $\alpha$ , *i.e.*, for which one does not find  $p_\mu \leq \alpha$ , constitute a confidence interval with confidence level  $1 - \alpha$ .

In general one will find that for some regions in the parameter space of the signal model, the predictions for data are almost indistinguishable from those of the background-only model. This corresponds to the case where  $\mu$  is very small, as would occur, *e.g.*, in a search for a new particle with a mass so high that its production rate in a given experiment is negligible. That is, one has essentially no experimental sensitivity to such a model.

One would prefer that if the sensitivity to a model (or a point in a model's parameter space) is very low, then it should not be excluded. Even if the outcomes predicted with or without signal are identical, however, the probability to reject the signal model will equal  $\alpha$ , the type-I error rate. As one often takes  $\alpha$  to be 5%, this would mean that in a large number of searches covering a broad range of a signal model's parameter space, there would inevitably be excluded regions in which the experimental sensitivity is very small, and thus one may question whether it is justified to regard such parameter values as disfavored.

Exclusion of models to which one has little or no sensitivity occurs, for example, if the data fluctuate very low relative to the expectation of the background-only hypothesis. In this case the resulting upper limit on  $\mu$  may be anomalously low. As a means of controlling this effect one often determines the mean or median limit under assumption of the background-only hypothesis, as discussed in Sec. 39.5.

One way to mitigate the problem of excluding models to which one is not sensitive is the  $\text{CL}_s$  method, where the measure used to test a parameter is increased for decreasing sensitivity [36,37]. The

procedure is based on a statistic called  $\text{CL}_s$ , which is defined as

$$\text{CL}_s = \frac{p_\mu}{1 - p_b}, \quad (39.73)$$

where  $p_b$  is the  $p$ -value of the background-only hypothesis. In the usual formulation of the method, both  $p_\mu$  and  $p_b$  are defined using a single test statistic, and the definition of  $\text{CL}_s$  above assumes this statistic is continuous; more details can be found in Refs. [36–37].

A point in a model's parameter space is regarded as excluded if one finds  $\text{CL}_s \leq \alpha$ . As the denominator in Eq. (39.73) is always less than or equal to unity, the exclusion criterion based on  $\text{CL}_s$  is more stringent than the usual requirement  $p_\mu \leq \alpha$ . In this sense the  $\text{CL}_s$  procedure is conservative, and the coverage probability of the corresponding intervals will exceed the nominal confidence level  $1 - \alpha$ . If the experimental sensitivity to a given value of  $\mu$  is very low, then one finds that as  $p_\mu$  decreases, so does the denominator  $1 - p_b$ , and thus the condition  $\text{CL}_s \leq \alpha$  is effectively prevented from being satisfied. In this way the exclusion of parameters in the case of low sensitivity is suppressed.

The  $\text{CL}_s$  procedure has the attractive feature that the resulting intervals coincide with those obtained from the Bayesian method in two important cases: the mean value of a Poisson or Gaussian distributed measurement with a constant prior. The  $\text{CL}_s$  intervals overcover for all values of the parameter  $\mu$ , however, by an amount that depends on  $\mu$ .

The problem of excluding parameter values to which one has little sensitivity is particularly acute when one wants to set a one-sided limit, *e.g.*, an upper limit on a cross section. Here one tests a value of a rate parameter  $\mu$  against the alternative of a lower rate, and therefore the critical region of the test is taken to correspond to data outcomes with a low event yield. If the number of events found in the search region fluctuates low enough, however, it can happen that all physically meaningful signal parameter values, including those to which one has very little sensitivity, are rejected by the test.

Another solution to this problem, therefore, is to replace the one-sided test by one based on the likelihood ratio, where the critical region is not restricted to low rates. This is the approach followed in the Feldman-Cousins procedure described in Section 39.4.2.1. The critical region for the test of a given value of  $\mu$  contains data values characteristic of both higher and lower rates. As a result, for a given observed rate one can in general obtain a two-sided interval. If, however, the parameter estimate  $\hat{\mu}$  is sufficiently close to the lower limit of zero, then only high values of  $\mu$  are rejected, and the lower edge of the confidence interval is at zero. Note, however, that the coverage property of  $1 - \alpha$  pertains to the entire interval, not to the probability for the upper edge  $\mu_{\text{up}}$  to be greater than the true value  $\mu$ . For parameter estimates increasingly far away from the boundary, *i.e.*, for increasing signal significance, the point  $\mu = 0$  is excluded and the interval has nonzero upper and lower edges.

An additional difficulty arises when a parameter estimate is not significantly far away from the boundary, in which case it is natural to report a one-sided confidence interval (often an upper limit). It is straightforward to force the Neyman prescription to produce only an upper limit by setting  $x_2 = \infty$  in Eq. (39.62). Then  $x_1$  is uniquely determined and the upper limit can be obtained. If, however, the data come out such that the parameter estimate is not so close to the boundary, one might wish to report a central confidence interval (*i.e.*, an interval based on a two-sided test with equal upper and lower tail areas). As pointed out by Feldman and Cousins [35], if the decision to report an upper limit or two-sided interval is made by looking at the data (“flip-flopping”), then in general there will be parameter values for which the resulting intervals have a coverage probability less than  $1 - \alpha$ . With the confidence intervals suggested in [35], the prescription determines whether the interval is one- or two-sided in a way which preserves the coverage probability (and are thus said to be *unified*).

The intervals according to this method for the mean of Poisson variable in the absence of background are given in Table 39.4. (Note that  $\alpha$  in Ref. 35 is defined following Neyman [31] as the coverage

probability; this is opposite the modern convention used here in which the coverage probability is  $1 - \alpha$ .) The values of  $1 - \alpha$  given here refer to the coverage of the true parameter by the whole interval  $[\mu_1, \mu_2]$ . In Table 39.3 for the one-sided upper limit, however,  $1 - \alpha$  refers to the probability to have  $\mu_{\text{up}} \geq \mu$  (or  $\mu_{\text{lo}} \leq \mu$  for lower limits).

**Table 39.4:** Unified confidence intervals  $[\mu_1, \mu_2]$  for a the mean of a Poisson variable given  $n$  observed events in the absence of background, for confidence levels of 90% and 95%.

$n$	$1 - \alpha = 90\%$		$1 - \alpha = 95\%$	
	$\mu_1$	$\mu_2$	$\mu_1$	$\mu_2$
0	0.00	2.44	0.00	3.09
1	0.11	4.36	0.05	5.14
2	0.53	5.91	0.36	6.72
3	1.10	7.42	0.82	8.25
4	1.47	8.60	1.37	9.76
5	1.84	9.99	1.84	11.26
6	2.21	11.47	2.21	12.75
7	3.56	12.53	2.58	13.81
8	3.96	13.99	2.94	15.29
9	4.36	15.30	4.36	16.77
10	5.50	16.50	4.75	17.82

A potential difficulty with unified intervals arises if, for example, one constructs such an interval for a Poisson parameter  $s$  of some yet to be discovered signal process with, say,  $1 - \alpha = 0.9$ . If the true signal parameter is zero, or in any case much less than the expected background, one will usually obtain a one-sided upper limit on  $s$ . In a certain fraction of the experiments, however, a two-sided interval for  $s$  will result. Since, however, one typically chooses  $1 - \alpha$  to be only 0.9 or 0.95 when setting limits, the value  $s = 0$  may be found below the lower edge of the interval before the existence of the effect is well established. It must then be communicated carefully that in excluding  $s = 0$  at, say, 90% or 95% confidence level from the interval, one is not necessarily claiming to have discovered the effect, for which one would usually require a higher level of significance (*e.g.*,  $5\sigma$ ).

Another possibility is to construct a Bayesian interval as described in Section 39.4.1. The presence of the boundary can be incorporated simply by setting the prior density to zero in the unphysical region. More specifically, the prior may be chosen using formal rules such as the reference prior or Jeffreys prior mentioned in Sec. 39.2.4.

In HEP a widely used prior for the mean  $\mu$  of a Poisson distributed measurement has been uniform for  $\mu \geq 0$ . This prior does not follow from any fundamental rule nor can it be regarded as reflecting a reasonable degree of belief, since the prior probability for  $\mu$  to lie between any two finite values is zero. The procedure above can be more appropriately regarded as a way for obtaining intervals with frequentist properties that can be investigated. The resulting upper limits have a coverage probability that depends on the true value of the Poisson parameter, and is nowhere smaller than the stated probability content. Lower limits and two-sided intervals for the Poisson mean based on flat priors undercover, however, for some values of the parameter, although to an extent that in practical cases may not be too severe [2,23]. Intervals constructed in this way have the advantage of being easy to derive; if several independent measurements are to be combined then one simply multiplies the likelihood functions (cf. Eq. (39.70)).

In any case, it is important to always report sufficient information so that the result can be combined with other measurements. Often this means giving an unbiased estimator and its standard deviation, even if the estimated value is in the unphysical region.

It can also be useful with a frequentist interval to calculate its subjective probability content using the posterior p.d.f. based on one or several reasonable guesses for the prior p.d.f. If it turns out to

be significantly less than the stated confidence level, this warns that it would be particularly misleading to draw conclusions about the parameter's value from the interval alone.

### 39.5. Experimental sensitivity

In this section we describe methods for characterizing the sensitivity of a search for a new physics signal. As discussed in Sec. 39.3, an experimental analysis can often be formulated as a test of hypothetical model parameters. Therefore we may quantify the sensitivity by giving the results that we expect from such a test under specific assumptions about the signal process.

Here to be concrete we will consider a parameter  $\mu$  proportional to the rate of a signal process, although the concepts described in this section may be easily generalized to other parameters. One may wish to establish discovery of the signal process by testing and rejecting the hypothesis that  $\mu = 0$ , and in addition one often wants to test nonzero values of  $\mu$  to construct a confidence interval (*e.g.*, limits) as described in Sec. 39.4. In the frequentist framework, the result of each tested value of  $\mu$  is the  $p$ -value  $p_\mu$  or equivalently the significance  $Z_\mu = \Phi^{-1}(1 - p_\mu)$ , where as usual  $\Phi$  is the standard Gaussian cumulative distribution and its inverse  $\Phi^{-1}$  is the standard Gaussian quantile.

Prior to carrying out the experiment, one generally wants to quantify what significance  $Z_\mu$  is expected under given assumptions for the presence or absence of the signal process. Specifically, for the significance of a test of  $\mu = 0$  (the discovery significance) one usually quotes the  $Z_0$  one would expect if the signal is present at a given nominal rate, which we can define in general to correspond to  $\mu = 1$ . For limits, one often gives the expected limit under assumption of the background-only ( $\mu = 0$ ) model. These quantities are used to optimize the analysis and to quantify the experimental sensitivity, that is, to characterize how likely it is to make a discovery if the signal is present, and to say what values of  $\mu$  one may be able to exclude if the signal is in fact absent.

First we clarify the notion of *expected significance*. Because the significance  $Z_\mu$  is a function of the data, it is itself a random quantity characterized by a certain sampling distribution. This distribution depends on the assumed value of  $\mu$ , which is not necessarily the same as the hypothesized value of  $\mu$  being tested. We may therefore consider the distribution  $f(Z_\mu|\mu')$ , *i.e.*, the distribution of  $Z_\mu$  that would be obtained by considering data samples generated under assumption of  $\mu'$ . In a similar way one can talk about the sampling distribution of an upper limit for  $\mu$ ,  $f(\mu_{\text{up}}|\mu')$ .

One can identify the expected significance or limit with either the mean or median of these distributions, but the median may be preferred since it is invariant under monotonic transformations. For example, the monotonic relation between  $p$ -value and significance,  $p = 1 - \Phi(Z)$ , then gives  $\text{med}[p_\mu|\mu'] = 1 - \Phi(\text{med}[Z_\mu|\mu'])$ , whereas the corresponding relation does not hold in general for the mean.

In some cases one may be able to write down approximate formulae for the distributions of  $Z_\mu$  and for limits, but more generally they must be determined from Monte Carlo calculations. In many cases of interest, the significance  $Z_\mu$  and the limits on  $\mu$  will have approximate Gaussian distributions.

As an example, consider a Poisson counting experiment, where the result consists of an observed number  $n$  of events, modeled as a Poisson distributed variable with a mean of  $\mu s + b$ . Here  $s$  and  $b$ , the expected numbers of events from signal and background processes, are taken to be known. If we are interested in discovering the signal process we test and try to reject the hypothesis  $\mu = 0$ . To characterize the experimental sensitivity, we want to give the discovery significance expected under the assumption of  $\mu = 1$ .

In the limit where its mean value is large, the Poisson variable  $n$  can be approximated as an almost continuous Gaussian variable with mean  $\mu s + b$  and standard deviation  $\sigma = \sqrt{\mu s + b}$ . In the usual case where a physical signal model corresponds to  $\mu > 0$ , the  $p$ -value of  $\mu = 0$  is the probability to find  $n$  greater than or equal to the value observed,

$$p_0 = \Phi\left(\frac{n - b}{\sqrt{b}}\right), \quad (39.74)$$

and the corresponding significance is  $Z_0 = \Phi^{-1}(1 - p_0) = (n - b)/\sqrt{b}$ . The median (here equal to the mean) of  $n$  assuming  $\mu = 1$  is  $s + b$ , and therefore the median discovery significance is

$$\text{med}[Z_0|\mu = 1] = \frac{s}{\sqrt{b}}. \quad (39.75)$$

The figure of merit “ $s/\sqrt{b}$ ” has been widely used in HEP as a measure of expected discovery significance. A better approximation for the Poisson counting experiment, however, may be obtained by testing  $\mu = 0$  using the likelihood ratio (39.43)  $\lambda(0) = L(0)/L(\hat{\mu})$ , where

$$L(\mu) = \frac{(\mu s + b)^n}{n!} e^{-(\mu s + b)} \quad (39.76)$$

is the likelihood function,  $\hat{\mu} = (n - b)/s$  is the ML estimator. In this example there are no nuisance parameters, as  $s$  and  $b$  are taken to be known. For the case where the relevant signal models correspond to positive  $\mu$ , one may test the  $\mu = 0$  hypothesis with the statistic  $q_0 = -2 \ln \lambda(0)$  when  $\hat{\mu} > 0$ , i.e., an excess is observed, and  $q_0 = 0$  otherwise. One can show (see, e.g., [40]) that in the large-sample limit, the discovery significance is then  $Z_0 = \sqrt{q_0}$ , for which one finds

$$Z_0 = \sqrt{2 \left( n \ln \frac{n}{b} + b - n \right)} \quad (39.77)$$

for  $n > b$  and  $Z_0 = 0$  otherwise. To approximate the expected discovery significance assuming  $\mu = 1$ , one may simply replace  $n$  with the expected value  $E[n|\mu = 1] = s + b$  (the so-called “Asimov data set”), giving

$$\text{med}[Z_0|\mu = 1] = \sqrt{2 \left( (s + b) \ln \left( 1 + \frac{s}{b} \right) - s \right)}. \quad (39.78)$$

This has been shown in Ref. 40 to provide a good approximation to the median discovery significance for values of  $s$  of several and for  $b$  well below unity. The right-hand side of Eq. (39.78) reduces to  $s/\sqrt{b}$  in the limit  $s \ll b$ .

Beyond the simple Poisson counting experiment, in general one may test values of a parameter  $\mu$  with more complicated functions of the measured data to obtain a  $p$ -value  $p_\mu$ , and from this one can quote the equivalent significance  $Z_\mu$  or find, e.g., an upper limit  $\mu_{\text{up}}$ . In this case as well one may quantify the experimental sensitivity by giving the significance  $Z_\mu$  expected if the data are generated with a different value of the parameter  $\mu'$ . In some problems, finding the sampling distribution of the significance or limits may be possible using large-sample formulae as described, e.g., in Ref. 40. In other cases a Monte Carlo study may be needed. Using whatever means available, one usually quotes the expected (mean or, preferably, median) significance or limit as the primary measures of experimental sensitivity.

Even if the true signal is present at its nominal rate, the actual discovery significance  $Z_0$  obtained from the real data is subject to statistical fluctuations and will not in general be equal to its expected value. In an analogous way, the observed limit will differ from the expected limit even if the signal is absent. Upon observing such a difference one would like to know how large this is compared to expected statistical fluctuations. Therefore, in addition to the observed significance and limits it is useful to communicate not only their expected values but also a measure of the width of their distributions.

As the distributions of significance and limits are often well approximated by a Gaussian, one may indicate the intervals corresponding to plus-or-minus one and/or two standard deviations. If the distributions are significantly non-Gaussian, one may use instead the quantiles that give the same probability content, i.e., [0.1587, 0.8413] for  $\pm 1\sigma$ , [0.02275, 0.97725] for  $\pm 2\sigma$ . An upper limit found significantly below the background-only expectation may indicate a strong downward fluctuation of the data, or perhaps as well an incorrect estimate of the background rate.

The procedures described above pertain to frequentist hypothesis tests and limits. Bayesian limits, just like those found from a

frequentist procedure, are functions of the data and one may therefore find, usually with approximations or Monte Carlo studies, their sampling distribution and corresponding mean (or, preferably, median) and standard deviation.

When trying to establish discovery of a signal process, the Bayesian approach may employ a Bayes factor as described in Sec. 39.3.3. In the case of the Poisson counting experiment with the likelihood from Eq. (39.76), the log of the Bayes factor that compares  $\mu = 1$  to  $\mu = 0$  is  $\ln B_{10} = \ln(L(1)/L(0)) = n \ln(1 + s/b) - s$ . That is, the expectation value, assuming  $\mu = 1$ , of  $\ln B_{10}$  for this problem is

$$E[\ln B_{10}|\mu = 1] = (s + b) \ln \left( 1 + \frac{s}{b} \right) - s. \quad (39.79)$$

Comparing this to Eq. (39.78), one finds  $\text{med}[Z_0|1] = \sqrt{2E[\ln B_{10}|1]}$ . Thus for this particular problem the frequentist median discovery significance can be related to the corresponding Bayes factor in a simple way.

In some analyses, the goal may not be to establish discovery of a signal process but rather to measure, as accurately as possible, the signal rate. If we consider again the Poisson counting experiment described by the likelihood function of Eq. (39.76), the ML estimator  $\hat{\mu} = (n - b)/s$  has a variance, assuming  $\mu = 1$ , of

$$V[\hat{\mu}] = V \left[ \frac{n - b}{s} \right] = \frac{1}{s^2} V[n] = \frac{s + b}{s^2}, \quad (39.80)$$

so that the standard deviation of  $\hat{\mu}$  is  $\sigma_{\hat{\mu}} = \sqrt{s + b}/s$ . One may therefore use  $s/\sqrt{s + b}$  as a figure of merit to be maximized in order to obtain the best measurement accuracy of a rate parameter. The quantity  $s/\sqrt{s + b}$  is also the expected significance with which one rejects  $s$  assuming the signal is absent, and thus can be used to optimize the expected upper limit on  $s$ .

## References:

1. B. Efron, Am. Stat. **40**, 11 (1986).
2. R.D. Cousins, Am. J. Phys. **63**, 398 (1995).
3. A. Stuart, J.K. Ord, and S. Arnold, *Kendall's Advanced Theory of Statistics*, Vol. 2A: *Classical Inference and the Linear Model*, 6th ed., Oxford Univ. Press (1999), and earlier editions by Kendall and Stuart. The likelihood-ratio ordering principle is described at the beginning of Ch. 23. Chapter 26 compares different schools of statistical inference.
4. F. James, *Statistical Methods in Experimental Physics*, 2nd ed., (World Scientific, Singapore, 2007).
5. H. Cramér, *Mathematical Methods of Statistics*, Princeton Univ. Press, New Jersey (1958).
6. L. Lyons, *Statistics for Nuclear and Particle Physicists*, (Cambridge University Press, New York, 1986).
7. R. Barlow, Nucl. Instrum. Methods **A297**, 496 (1990).
8. G. Cowan, *Statistical Data Analysis*, (Oxford University Press, Oxford, 1998).
9. For a review, see S. Baker and R. Cousins, Nucl. Instrum. Methods **221**, 437 (1984).
10. S.S. Wilks, Ann. Math. Statist. **9**, 60 (1938).
11. Christopher M. Bishop, *Pattern Recognition and Machine Learning*, (Springer, New York, 2006).
12. T. Hastie, R. Tibshirani, and J. Friedman, *The Elements of Statistical Learning*, (2nd edition, Springer, New York, 2009).
13. A. Webb, *Statistical Pattern Recognition*, 2nd ed., (Wiley, New York, 2002).
14. L.I. Kuncheva, *Combining Pattern Classifiers*, (Wiley, New York, 2004).
15. Links to the *Proceedings of the PHYSTAT* conference series (Durham 2002, Stanford 2003, Oxford 2005, and Geneva 2007, 2011) can be found at [phystat.org](http://phystat.org).
16. A. Höcker et al., *TMVA Users Guide*, physics/0703039(2007); software available from [tmva.sourceforge.net](http://tmva.sourceforge.net).
17. I. Narsky, *StatPatternRecognition: A C++ Package for Statistical Analysis of High Energy Physics Data*, physics/0507143(2005); software avail. from [sourceforge.net/projects/statpatrec](http://sourceforge.net/projects/statpatrec).

18. F. Pedregosa et al., *Scikit-learn: Machine Learning in Python*, Journal of Machine Learning Research **12** 2825–2830 (2011).
19. L. Demortier, *P-Values and Nuisance Parameters, Proceedings of PHYSTAT 2007*, CERN-2008-001, p. 23.
20. E. Gross and O. Vitells, Eur. Phys. J. **C70**, 525 (2010); [arXiv:1005.1891](#).
21. R.B. Davis, Biometrika **74**, 33 (1987).
22. K. Cranmer, *Statistical Challenges for Searches for New Physics at the LHC*, in Proceedings of PHYSTAT 2005, L. Lyons and M. Karagoz Unel (eds.), Oxford (2005); [arXiv:physics/0511028](#).
23. B.P. Roe and M.B. Woodroffe, Phys. Rev. **D63**, 13009 (2000).
24. A. O'Hagan and J.J. Forster, *Bayesian Inference*, (2nd edition, volume 2B of *Kendall's Advanced Theory of Statistics*, Arnold, London, 2004).
25. D. Sivia and J. Skilling, *Data Analysis: A Bayesian Tutorial*, (Oxford University Press, 2006).
26. P.C. Gregory, *Bayesian Logical Data Analysis for the Physical Sciences*, (Cambridge University Press, 2005).
27. J.M. Bernardo and A.F.M. Smith, *Bayesian Theory*, (Wiley, 2000).
28. Robert E. Kass and Larry Wasserman, J. Am. Stat. Assoc. **91**, 1343 (1996).
29. J.M. Bernardo, J. R. Statist. Soc. **B41**, 113 (1979); J.M. Bernardo and J.O. Berger, J. Am. Stat. Assoc. **84**, 200 (1989). See also J.M. Bernardo, *Reference Analysis*, in *Handbook of Statistics*, 25 (D.K. Dey and C.R. Rao, eds.), 17-90, Elsevier (2005) and references therein.
30. L. Demortier, S. Jain, and H. Prosper, Phys. Rev. **D82**, 034002 (2010) [[arXiv:1002.1111](#)].
31. J. Neyman, Phil. Trans. Royal Soc. London, Series A, **236**, 333 (1937), reprinted in *A Selection of Early Statistical Papers on J. Neyman*, (University of California Press, Berkeley, 1967).
32. R. E. Kass and A. E. Raftery, J. Am. Stat. Assoc. **90**, 773 (1995).
33. J. Skilling, *Nested Sampling, AIP Conference Proceedings*, **735**, 395–405 (2004).
34. F. Feroz, M.P. Hobson, and M. Bridges, Mon. Not. Roy. Astron. Soc. **398**, 1601-1614 (2009); [arXiv:0809.3437](#).
35. G.J. Feldman and R.D. Cousins, Phys. Rev. **D57**, 3873 (1998). This paper does not specify what to do if the ordering principle gives equal rank to some values of  $x$ . Eq. 21.6 of Ref. [3] gives the rule: all such points are included in the acceptance region (the domain  $D(\alpha)$ ). Some authors have assumed the contrary, and shown that one can then obtain null intervals.
36. A.L. Read, *Modified frequentist analysis of search results (the  $CL_s$  method)*, in F. James, L. Lyons, and Y. Perrin (eds.), *Workshop on Confidence Limits*, CERN Yellow Report 2000-005, available through [cdsweb.cern.ch](#).
37. T. Junk, Nucl. Instrum. Methods **A434**, 435 (1999).
38. N. Reid, *Likelihood Inference in the Presence of Nuisance Parameters, Proceedings of PHYSTAT2003*, L. Lyons, R. Mount, and R. Reitmeyer, eds., eConf C030908, Stanford, 2003.
39. W.A. Rolke, A.M. Lopez, and J. Conrad, Nucl. Instrum. Methods **A551**, 493 (2005).
40. G. Cowan et al., Eur. Phys. J. **C71**, 1554 (2011).



## 40. MONTE CARLO TECHNIQUES

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Monte Carlo techniques are often the only practical way to evaluate difficult integrals or to sample random variables governed by complicated probability density functions. Here we describe an assortment of methods for sampling some commonly occurring probability density functions.

### 40.1. Sampling the uniform distribution

Most Monte Carlo sampling or integration techniques assume a “random number generator,” which generates uniform statistically independent values on the half open interval  $[0, 1)$ ; for reviews see, *e.g.*, [1,2].

Uniform random number generators are available in software libraries such as CERNLIB [3], CLHEP [4], and ROOT [5]. For example, in addition to a basic congruential generator `TRandom` (see below), ROOT provides three more sophisticated routines: `TRandom1` implements the RANLUX generator [6] based on the method by Lüscher, and allows the user to select different quality levels, trading off quality with speed; `TRandom2` is based on the maximally equidistributed combined Tausworthe generator by L’Ecuyer [7]; the `TRandom3` generator implements the Mersenne twister algorithm of Matsumoto and Nishimura [8]. All of the algorithms produce a periodic sequence of numbers, and to obtain effectively random values, one must not use more than a small subset of a single period. The Mersenne twister algorithm has an extremely long period of  $2^{19937} - 1$ .

The performance of the generators can be investigated with tests such as DIEHARD [9] or TestU01 [10]. Many commonly available congruential generators fail these tests and often have sequences (typically with periods less than  $2^{32}$ ), which can be easily exhausted on modern computers. A short period is a problem for the `TRandom` generator in ROOT, which, however, has the advantage that its state is stored in a single 32-bit word. The generators `TRandom1`, `TRandom2`, or `TRandom3` have much longer periods, with `TRandom3` being recommended by the ROOT authors as providing the best combination of speed and good random properties. For further information see, *e.g.*, Ref. 11.

### 40.2. Inverse transform method

If the desired probability density function is  $f(x)$  on the range  $-\infty < x < \infty$ , its cumulative distribution function (expressing the probability that  $x \leq a$ ) is given by Eq. (38.6). If  $a$  is chosen with probability density function  $f(a)$ , then the integrated probability up to point  $a$ ,  $F(a)$ , is itself a random variable which will occur with uniform probability density on  $[0, 1]$ . Suppose  $u$  is generated according to a uniformly distributed in  $(0, 1)$ . If  $x$  can take on any value, and ignoring the endpoints, we can then find a unique  $x$  chosen from the p.d.f.  $f(x)$  for a given  $u$  if we set

$$u = F(x), \quad (40.1)$$

provided we can find an inverse of  $F$ , defined by

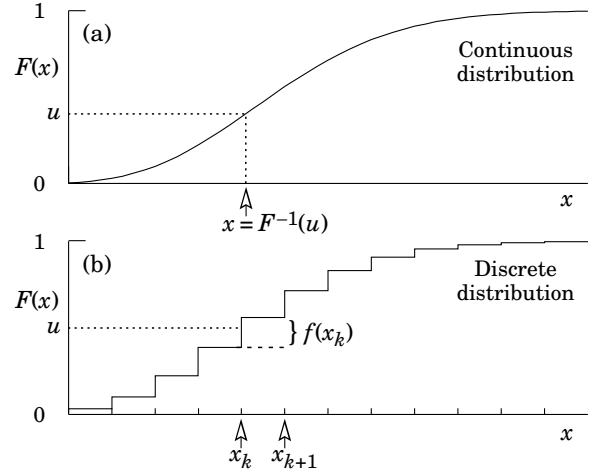
$$x = F^{-1}(u). \quad (40.2)$$

This method is shown in Fig. 40.1a. It is most convenient when one can calculate by hand the inverse function of the indefinite integral of  $f$ . This is the case for some common functions  $f(x)$  such as  $\exp(x)$ ,  $(1-x)^n$ , and  $1/(1+x^2)$  (Cauchy or Breit-Wigner), although it does not necessarily produce the fastest generator. Standard libraries contain software to implement this method numerically, working from functions or histograms in one or more dimensions, *e.g.*, the UNU.RAN package [12], available in ROOT.

For a discrete distribution,  $F(x)$  will have a discontinuous jump of size  $f(x_k)$  at each allowed  $x_k, k = 1, 2, \dots$ . Choose  $u$  from a uniform distribution on  $(0, 1)$  as before. Find  $x_k$  such that

$$F(x_{k-1}) < u \leq F(x_k) \equiv \text{Prob}(x \leq x_k) = \sum_{i=1}^k f(x_i); \quad (40.3)$$

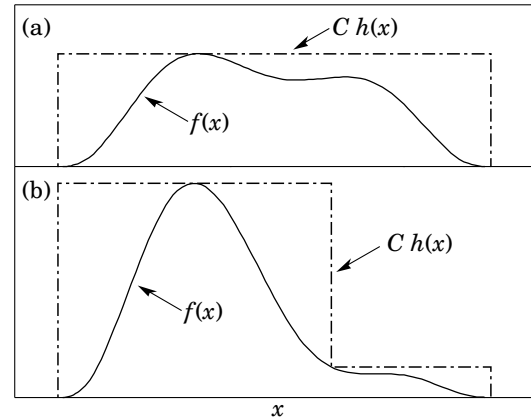
then  $x_k$  is the value we seek (note:  $F(x_0) \equiv 0$ ). This algorithm is illustrated in Fig. 40.1b.



**Figure 40.1:** Use of a random number  $u$  chosen from a uniform distribution  $(0,1)$  to find a random number  $x$  from a distribution with cumulative distribution function  $F(x)$ .

### 40.3. Acceptance-rejection method (Von Neumann)

Very commonly an analytic form for  $F(x)$  is unknown or too complex to work with, so that obtaining an inverse as in Eq. (40.2) is impractical. We suppose that for any given value of  $x$ , the probability density function  $f(x)$  can be computed, and further that enough is known about  $f(x)$  that we can enclose it entirely inside a shape which is  $C$  times an easily generated distribution  $h(x)$ , as illustrated in Fig. 40.2. That is,  $Ch(x) \geq f(x)$  must hold for all  $x$ .



**Figure 40.2:** Illustration of the acceptance-rejection method. Random points are chosen inside the upper bounding figure, and rejected if the ordinate exceeds  $f(x)$ . The lower figure illustrates a method to increase the efficiency (see text).

Frequently  $h(x)$  is uniform or is a normalized sum of uniform distributions. Note that both  $f(x)$  and  $h(x)$  must be normalized to unit area, and therefore, the proportionality constant  $C > 1$ . To generate  $f(x)$ , first generate a candidate  $x$  according to  $h(x)$ . Calculate  $f(x)$  and the height of the envelope  $Ch(x)$ ; generate  $u$  and test if  $uC h(x) \leq f(x)$ . If so, accept  $x$ ; if not reject  $x$  and try again. If we regard  $x$  and  $uC h(x)$  as the abscissa and ordinate of a point in a two-dimensional plot, these points will populate the entire area  $Ch(x)$  in a smooth manner; then we accept those which fall under  $f(x)$ . The efficiency is the ratio of areas, which must equal  $1/C$ ; therefore we must keep  $C$  as close as possible to 1.0. Therefore, we try to choose  $Ch(x)$  to be as close to  $f(x)$  as convenience dictates, as in the lower part of Fig. 40.2.

#### 40.4. Algorithms

Algorithms for generating random numbers belonging to many different distributions are given for example by Press [13], Ahrens and Dieter [14], Rubinstein [15], Devroye [16], Walck [17] and Gentle [18]. For many distributions, alternative algorithms exist, varying in complexity, speed, and accuracy. For time-critical applications, these algorithms may be coded in-line to remove the significant overhead often encountered in making function calls.

In the examples given below, we use the notation for the variables and parameters given in Table 38.1. Variables named “ $u$ ” are assumed to be independent and uniform on  $[0,1]$ . Denominators must be verified to be non-zero where relevant.

##### 40.4.1. Exponential decay :

This is a common application of the inverse transform method, and uses the fact that if  $u$  is uniformly distributed in  $[0,1]$ , then  $(1-u)$  is as well. Consider an exponential p.d.f.  $f(t) = (1/\tau) \exp(-t/\tau)$  that is truncated so as to lie between two values,  $a$  and  $b$ , and renormalized to unit area. To generate decay times  $t$  according to this p.d.f., first let  $\alpha = \exp(-a/\tau)$  and  $\beta = \exp(-b/\tau)$ ; then generate  $u$  and let

$$t = -\tau \ln(\beta + u(\alpha - \beta)). \quad (40.4)$$

For  $(a, b) = (0, \infty)$ , we have simply  $t = -\tau \ln u$ . (See also Sec. 40.4.6.)

##### 40.4.2. Isotropic direction in 3D :

Isotropy means the density is proportional to solid angle, the differential element of which is  $d\Omega = d(\cos\theta)d\phi$ . Hence  $\cos\theta$  is uniform  $(2u_1 - 1)$  and  $\phi$  is uniform  $(2\pi u_2)$ . For alternative generation of  $\sin\phi$  and  $\cos\phi$ , see the next subsection.

##### 40.4.3. Sine and cosine of random angle in 2D :

Generate  $u_1$  and  $u_2$ . Then  $v_1 = 2u_1 - 1$  is uniform on  $(-1,1)$ , and  $v_2 = u_2$  is uniform on  $(0,1)$ . Calculate  $r^2 = v_1^2 + v_2^2$ . If  $r^2 > 1$ , start over. Otherwise, the sine ( $S$ ) and cosine ( $C$ ) of a random angle (*i.e.*, uniformly distributed between zero and  $2\pi$ ) are given by

$$S = 2v_1v_2/r^2 \quad \text{and} \quad C = (v_1^2 - v_2^2)/r^2. \quad (40.5)$$

##### 40.4.4. Gaussian distribution :

If  $u_1$  and  $u_2$  are uniform on  $(0,1)$ , then

$$z_1 = \sin(2\pi u_1) \sqrt{-2 \ln u_2} \quad \text{and} \quad z_2 = \cos(2\pi u_1) \sqrt{-2 \ln u_2} \quad (40.6)$$

are independent and Gaussian distributed with mean 0 and  $\sigma = 1$ .

There are many variants of this basic algorithm, which may be faster. For example, construct  $v_1 = 2u_1 - 1$  and  $v_2 = 2u_2 - 1$ , which are uniform on  $(-1,1)$ . Calculate  $r^2 = v_1^2 + v_2^2$ , and if  $r^2 > 1$  start over. If  $r^2 < 1$ , it is uniform on  $(0,1)$ . Then

$$z_1 = v_1 \sqrt{\frac{-2 \ln r^2}{r^2}} \quad \text{and} \quad z_2 = v_2 \sqrt{\frac{-2 \ln r^2}{r^2}} \quad (40.7)$$

are independent numbers chosen from a normal distribution with mean 0 and variance 1.  $z'_i = \mu + \sigma z_i$  distributes with mean  $\mu$  and variance  $\sigma^2$ .

For a multivariate Gaussian with an  $n \times n$  covariance matrix  $V$ , one can start by generating  $n$  independent Gaussian variables,  $\{\eta_j\}$ , with mean 0 and variance 1 as above. Then the new set  $\{x_i\}$  is obtained as  $x_i = \mu_i + \sum_j L_{ij} \eta_j$ , where  $\mu_i$  is the mean of  $x_i$ , and  $L_{ij}$  are the components of  $L$ , the unique lower triangular matrix that fulfils  $V = LL^T$ . The matrix  $L$  can be easily computed by the following recursive relation (Cholesky's method):

$$L_{jj} = \left( V_{jj} - \sum_{k=1}^{j-1} L_{jk}^2 \right)^{1/2}, \quad (40.8a)$$

$$L_{ij} = \frac{V_{ij} - \sum_{k=1}^{j-1} L_{ik} L_{jk}}{L_{jj}}, \quad j = 1, \dots, n; \quad i = j+1, \dots, n, \quad (40.8b)$$

where  $V_{ij} = \rho_{ij}\sigma_i\sigma_j$  are the components of  $V$ . For  $n = 2$  one has

$$L = \begin{pmatrix} \sigma_1 & 0 \\ \rho\sigma_2 & \sqrt{1 - \rho^2}\sigma_2 \end{pmatrix}, \quad (40.9)$$

and therefore the correlated Gaussian variables are generated as  $x_1 = \mu_1 + \sigma_1\eta_1$ ,  $x_2 = \mu_2 + \rho\sigma_2\eta_1 + \sqrt{1 - \rho^2}\sigma_2\eta_2$ .

##### 40.4.5. $\chi^2(n)$ distribution :

To generate a variable following the  $\chi^2$  distribution for  $n$  degrees of freedom, use the Gamma distribution with  $k = n/2$  and  $\lambda = 1/2$  using the method of Sec. 40.4.6.

##### 40.4.6. Gamma distribution :

All of the following algorithms are given for  $\lambda = 1$ . For  $\lambda \neq 1$ , divide the resulting random number  $x$  by  $\lambda$ .

- If  $k = 1$  (the *exponential* distribution), accept  $x = -\ln u$ . (See also Sec. 40.4.1.)
- If  $0 < k < 1$ , initialize with  $v_1 = (e + k)/e$  (with  $e = 2.71828\dots$  being the natural log base). Generate  $u_1, u_2$ . Define  $v_2 = v_1 u_1$ .

**Case 1:**  $v_2 \leq 1$ . Define  $x = v_2^{1/k}$ . If  $u_2 \leq e^{-x}$ , accept  $x$  and stop, else restart by generating new  $u_1, u_2$ .

**Case 2:**  $v_2 > 1$ . Define  $x = -\ln([v_1 - v_2]/k)$ . If  $u_2 \leq x^{k-1}$ , accept  $x$  and stop, else restart by generating new  $u_1, u_2$ . Note that, for  $k < 1$ , the probability density has a pole at  $x = 0$ , so that return values of zero due to underflow must be accepted or otherwise dealt with.

- Otherwise, if  $k > 1$ , initialize with  $c = 3k - 0.75$ . Generate  $u_1$  and compute  $v_1 = u_1(1 - u_1)$  and  $v_2 = (u_1 - 0.5)\sqrt{c/v_1}$ . If  $x = k + v_2 - 1 \leq 0$ , go back and generate new  $u_1$ ; otherwise generate  $u_2$  and compute  $v_3 = 64v_1^3 u_2^2$ . If  $v_3 \leq 1 - 2v_2^2/x$  or if  $\ln v_3 \leq 2\{[k-1] \ln[x/(k-1)] - v_2\}$ , accept  $x$  and stop; otherwise go back and generate new  $u_1$ .

##### 40.4.7. Binomial distribution :

Begin with  $k = 0$  and generate  $u$  uniform in  $[0,1]$ . Compute  $P_k = (1-p)^n$  and store  $P_k$  into  $B$ . If  $u \leq B$  accept  $r_k = k$  and stop. Otherwise, increment  $k$  by one; compute the next  $P_k$  as  $P_k \cdot (p/(1-p)) \cdot (n-k)/(k+1)$ ; add this to  $B$ . Again, if  $u \leq B$ , accept  $r_k = k$  and stop, otherwise iterate until a value is accepted. If  $p > 1/2$ , it will be more efficient to generate  $r$  from  $f(r; n, q)$ , *i.e.*, with  $p$  and  $q$  interchanged, and then set  $r_k = n - r$ .

##### 40.4.8. Poisson distribution :

Iterate until a successful choice is made: Begin with  $k = 1$  and set  $A = 1$  to start. Generate  $u$ . Replace  $A$  with  $uA$ ; if now  $A < \exp(-\mu)$ , where  $\mu$  is the Poisson parameter, accept  $n_k = k - 1$  and stop. Otherwise increment  $k$  by 1, generate a new  $u$  and repeat, always starting with the value of  $A$  left from the previous try.

Note that the Poisson generator used in ROOT's `TRandom` classes before version 5.12 (including the derived classes `TRandom1`, `TRandom2`, `TRandom3`) as well as the routine `RNPSSN` from CERNLIB, use a Gaussian approximation when  $\mu$  exceeds a given threshold. This may be satisfactory (and much faster) for some applications. To do this, generate  $z$  from a Gaussian with zero mean and unit standard deviation; then use  $x = \max(0, [\mu + z\sqrt{\mu} + 0.5])$  where  $[ ]$  signifies the greatest integer  $\leq$  the expression. The routines from Numerical Recipes [13] and CLHEP's routine `RandPoisson` do not make this approximation (see, *e.g.*, Ref. 11).

##### 40.4.9. Student's $t$ distribution :

Generate  $u_1$  and  $u_2$  uniform in  $(0,1)$ ; then  $t = \sin(2\pi u_1)[n(u_2^{-2/n} - 1)]^{1/2}$  follows the Student's  $t$  distribution for  $n > 0$  degrees of freedom ( $n$  not necessarily an integer).

Alternatively, generate  $x$  from a Gaussian with mean 0 and  $\sigma^2 = 1$  according to the method of 40.4.4. Next generate  $y$ , an independent gamma random variate, according to 40.4.6 with  $\lambda = 1/2$  and  $k = n/2$ . Then  $z = x/\sqrt{y/n}$  is distributed as a  $t$  with  $n$  degrees of freedom.

For the special case  $n = 1$ , the Breit-Wigner distribution, generate  $u_1$  and  $u_2$ ; set  $v_1 = 2u_1 - 1$  and  $v_2 = 2u_2 - 1$ . If  $v_1^2 + v_2^2 \leq 1$  accept  $z = v_1/v_2$  as a Breit-Wigner distribution with unit area, center at 0.0, and FWHM 2.0. Otherwise start over. For center  $M_0$  and FWHM  $\Gamma$ , use  $W = z\Gamma/2 + M_0$ .

#### 40.4.10. Beta distribution :

The choice of an appropriate algorithm for generation of beta distributed random numbers depends on the values of the parameters  $\alpha$  and  $\beta$ . For, e.g.,  $\alpha = 1$ , one can use the transformation method to find  $x = 1 - u^{1/\beta}$ , and similarly if  $\beta = 1$  one has  $x = u^{1/\alpha}$ . For more general cases see, e.g., Refs. [17,18] and references therein.

### 40.5. Markov Chain Monte Carlo

In applications involving generation of random numbers following a multivariate distribution with a high number of dimensions, the transformation method may not be possible and the acceptance-rejection technique may have too low of an efficiency to be practical. If it is not required to have independent random values, but only that they follow a certain distribution, then Markov Chain Monte Carlo (MCMC) methods can be used. In depth treatments of MCMC can be found, e.g., in the texts by Robert and Casella [19], Liu [20], and the review by Neal [21].

MCMC is particularly useful in connection with Bayesian statistics, where a p.d.f.  $p(\theta)$  for an  $n$ -dimensional vector of parameters  $\theta = (\theta_1, \dots, \theta_n)$  is obtained, and one needs the marginal distribution of a subset of the components. Here one samples  $\theta$  from  $p(\theta)$  and simply records the marginal distribution for the components of interest.

A simple and broadly applicable MCMC method is the Metropolis-Hastings algorithm, which allows one to generate multidimensional points  $\theta$  distributed according to a target p.d.f. that is proportional to a given function  $p(\theta)$ . It is not necessary to have  $p(\theta)$  normalized to unit area, which is useful in Bayesian statistics, as posterior probability densities are often determined only up to an unknown normalization constant.

To generate points that follow  $p(\theta)$ , one first needs a proposal p.d.f.  $q(\theta; \theta_0)$ , which can be (almost) any p.d.f. from which independent random values  $\theta$  can be generated, and which contains as a parameter another point in the same space  $\theta_0$ . For example, a multivariate Gaussian centered about  $\theta_0$  can be used. Beginning at an arbitrary starting point  $\theta_0$ , the Hastings algorithm iterates the following steps:

1. Generate a value  $\theta$  using the proposal density  $q(\theta; \theta_0)$ ;
2. Form the Hastings test ratio,  $\alpha = \min \left[ 1, \frac{p(\theta)q(\theta_0; \theta)}{p(\theta_0)q(\theta; \theta_0)} \right]$ ;
3. Generate a value  $u$  uniformly distributed in  $[0, 1]$ ;
4. If  $u \leq \alpha$ , take  $\theta_1 = \theta$ . Otherwise, repeat the old point, i.e.,  $\theta_1 = \theta_0$ .
5. Set  $\theta_0 = \theta_1$  and return to step 1.

If one takes the proposal density to be symmetric in  $\theta$  and  $\theta_0$ , then this is the *Metropolis-Hastings* algorithm, and the test ratio becomes  $\alpha = \min[1, p(\theta)/p(\theta_0)]$ . That is, if the proposed  $\theta$  is at a value of probability higher than  $\theta_0$ , the step is taken. If the proposed step is rejected, the old point is repeated.

Methods for assessing and optimizing the performance of the algorithm are discussed in, e.g., Refs. [19–21]. One can, for example, examine the autocorrelation as a function of the lag  $k$ , i.e., the correlation of a sampled point with that  $k$  steps removed. This should decrease as quickly as possible for increasing  $k$ .

Generally one chooses the proposal density so as to optimize some quality measure such as the autocorrelation. For certain problems it has been shown that one achieves optimal performance when the acceptance fraction, that is, the fraction of points with  $u \leq \alpha$ , is around 40%. This can be adjusted by varying the width of the proposal density. For example, one can use for the proposal p.d.f. a multivariate Gaussian with the same covariance matrix as that of the target p.d.f., but scaled by a constant.

#### References:

1. F. James, *Comp. Phys. Comm.* **60**, 329 (1990).
2. P. L'Ecuyer, *Proc. 1997 Winter Simulation Conference*, IEEE Press, Dec. 1997, 127–134.
3. The CERN Program Library (CERNLIB); see [cernlib.web.cern.ch/cernlib](http://cernlib.web.cern.ch/cernlib).
4. Leif Lönnblad, *Comp. Phys. Comm.* **84**, 307 (1994).
5. Rene Brun and Fons Rademakers, *Nucl. Inst. Meth.* **A389**, 81 (1997); see also [root.cern.ch](http://root.cern.ch).
6. F. James, *Comp. Phys. Comm.* **79**, 111 (1994), based on M. Lüscher, *Comp. Phys. Comm.* **79**, 100 (1994).
7. P. L'Ecuyer, *Mathematics of Computation*, **65**, 213 (1996) and **65**, 225 (1999).
8. M. Matsumoto and T. Nishimura, *ACM Transactions on Modeling and Computer Simulation*, Vol. 8, No. 1, January 1998, 3–30.
9. Much of DIEHARD is described in: G. Marsaglia, *A Current View of Random Number Generators*, keynote address, *Computer Science and Statistics: 16th Symposium on the Interface*, Elsevier (1985).
10. P. L'Ecuyer and R. Simard, *ACM Transactions on Mathematical Software* **33**, 4, Article 1, December 2007.
11. J. Heinrich, CDF Note CDF/MEMO/STATISTICS/PUBLIC/8032, 2006.
12. UNU.RAN is described at [statmath.wu.ac.at/software/unuran](http://statmath.wu.ac.at/software/unuran); see also W. Hörmann, J. Leydold, and G. Derflinger, *Automatic Nonuniform Random Variate Generation*, (Springer, New York, 2004).
13. W.H. Press *et al.*, *Numerical Recipes*, 3rd edition, (Cambridge University Press, New York, 2007).
14. J.H. Ahrens and U. Dieter, *Computing* **12**, 223 (1974).
15. R.Y. Rubinstein, *Simulation and the Monte Carlo Method*, (John Wiley and Sons, Inc., New York, 1981).
16. L. Devroye, *Non-Uniform Random Variate Generation*, (Springer-Verlag, New York, 1986); available online at [cg.scs.carleton.ca/~luc/rnbookindex.html](http://cg.scs.carleton.ca/~luc/rnbookindex.html).
17. C. Walck, *Handbook on Statistical Distributions for Experimentalists*, University of Stockholm Internal Report SUF-PFY/96-01, available from [www.physto.se/~walck](http://www.physto.se/~walck).
18. J.E. Gentle, *Random Number Generation and Monte Carlo Methods*, 2nd ed., (Springer, New York, 2003).
19. C.P. Robert and G. Casella, *Monte Carlo Statistical Methods*, 2nd ed., (Springer, New York, 2004).
20. J.S. Liu, *Monte Carlo Strategies in Scientific Computing*, (Springer, New York, 2001).
21. R.M. Neal, *Probabilistic Inference Using Markov Chain Monte Carlo Methods*, Technical Report CRG-TR-93-1, Dept. of Computer Science, University of Toronto, available from [www.cs.toronto.edu/~radford/res-mcmc.html](http://www.cs.toronto.edu/~radford/res-mcmc.html).

## 41. MONTE CARLO EVENT GENERATORS

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General-purpose Monte Carlo (GPMC) generators like HERWIG [1], HERWIG++ [2], PYTHIA 6 [3], PYTHIA 8 [4], and SHERPA [5], provide fully exclusive simulations of high-energy collisions. They play an essential role in QCD modeling (in particular for aspects beyond fixed-order perturbative QCD), in data analysis, where they are used together with detector simulation to provide a realistic estimate of the detector response to collision events, and in the planning of new experiments, where they are used to estimate signals and backgrounds in high-energy processes. They are built from several components, that describe the physics starting from very short distance scales, up to the typical scale of hadron formation and decay. Since QCD is weakly interacting at short distances (below a femtometer), the components of the GPMC dealing with short-distance physics are based upon perturbation theory. At larger distances, all soft hadronic phenomena, like hadronization and the formation of the underlying event, cannot be computed from first principles, and one must rely upon QCD-inspired models.

The purpose of this review is to illustrate the main components of these generators. It is divided into four sections. The first one deals with short-distance, perturbative phenomena. The basic concepts leading to the simulations of the dominant QCD processes are illustrated here. In the second section, hadronization phenomena are treated. The two most popular hadronization models for the formation of primary hadrons, the string and cluster models, are illustrated. The basics of the implementation of primary-hadron decays into stable ones is also illustrated here. In the third section, models for soft hadron physics are discussed. These include models for the underlying event, and for minimum-bias interactions. Issues of Bose-Einstein and color-reconnection effects are also discussed here. The fourth section briefly introduces the problem of MC tuning.

We use natural units throughout, such that  $c = 1$  and  $\hbar = 1$ , with energy, momenta and masses measured in GeV, and time and distances measured in  $\text{GeV}^{-1}$ .

### 41.1. Short-distance physics in GPMC generators

The short-distance components of a GPMC generator deal with the computation of the primary process at hand, with decays of short-lived particles, and with the generation of QCD and QED radiation, on time scales below  $1/\Lambda$ , with  $\Lambda$  denoting a typical hadronic scale of a few hundred MeV, corresponding roughly to an inverse femtometer. In  $e^+e^-$  annihilation, for example, the short-distance physics describes the evolution of the system from the instant when the  $e^+e^-$  pair annihilates up to a time when the size of the produced system is just below a femtometer.

In the present discussion we take the momentum scale of the primary process to be  $Q \gg \Lambda$ , so that the corresponding time and distance scale  $1/Q$  is small. Soft- and collinear-safe inclusive observables, such as total decay widths or inclusive cross sections, can then be reliably computed in QCD perturbation theory (pQCD), with the perturbative expansion truncated at any fixed order  $n$ , and the remainder suppressed by  $\alpha_S(Q)^{n+1}$ .

Less inclusive observables, however, can receive large enhancements that destroy the convergence of the fixed-order expansion. This is due to the presence of collinear and infrared singularities in QCD. Thus, for example, a correction in which a parton from the primary interaction splits collinearly into two partons of comparable energy, is of order  $\alpha_S(Q) \ln(Q/\Lambda)$ , where the logarithm arises from an integral over a singularity regulated by the hadronic scale  $\Lambda$ . Since  $\alpha_S(Q) \propto 1/\ln(Q/\Lambda)$ , the corresponding cross section receives a correction of order unity. Two subsequent collinear splittings yield  $\alpha_S^2(Q) \ln^2(Q/\Lambda)$ , and so on. Thus, corrections of order unity arise at all orders in perturbation theory. The dominant region of phase space is the one where radiation is strongly ordered in a measure of hardness. This means that, from a typical final-state configuration, by clustering together final-state parton pairs with the smallest hardness recursively, we can reconstruct a branching tree, that may be viewed as the splitting history of the event. This history necessarily has some dependence on how we define hardness. For example, we can define it

as the energy of the incoming parton times the splitting angle, or as its virtuality, or as the transverse momentum of the splitting partons with respect to the incoming one. These definitions, however, are all equivalent in the collinear region. In fact, in the small-angle limit, the virtuality of a parton of energy  $E$ , splitting into two on-shell partons, is given by

$$p^2 = E^2 z(1-z)(1-\cos\theta) \approx \frac{z(1-z)}{2} E^2 \theta^2, \quad (41.1)$$

where  $z$  and  $1-z$  are the energy fractions carried by the produced partons, and  $\theta$  is their relative angle. The transverse momentum of the final partons relative to the direction of the incoming one is given by

$$p_T^2 \approx z^2(1-z)^2 E^2 \theta^2. \quad (41.2)$$

Thus, significant differences between these measures only arise in regions with very small  $z$  or  $1-z$  values. In QCD, because of soft divergences, these regions are in fact important, and the choice of the appropriate ordering variable is very relevant (see Sec. 41.3).

The so called KLN theorem [6,7] guarantees that large logarithmically divergent corrections, arising from final-state collinear splitting and from soft emissions, cancel against the virtual corrections in the total cross section, order by order in perturbation theory. Furthermore, the factorization theorem guarantees that initial-state collinear singularities can be factorized into the parton density functions (PDFs). Therefore, the cross section for the basic process remains accurate up to corrections of higher orders in  $\alpha_S(Q)$ , provided it is interpreted as an inclusive cross section, rather than as a bare partonic cross section. Thus, for example, the leading order (LO) cross section for  $e^+e^- \rightarrow q\bar{q}$  is a good LO estimate of the  $e^+e^-$  cross section for the production of a pair of quarks accompanied by an arbitrary number of collinear and soft gluons, but is not a good estimate of the cross section for the production of a  $q\bar{q}$  pair with no extra radiation.

Shower algorithms are used to compute the cross section for generic hard processes including all leading-logarithmic (LL) corrections. These algorithms begin with the generation of the kinematics of the basic process, performed with a probability proportional to its LO partonic cross section, which is interpreted physically as the inclusive cross section for the basic process, followed by an arbitrary sequence of shower splittings. A probability is then assigned to each splitting sequence. Thus, the initial LO cross section is partitioned into the cross sections for a multitude of final states of arbitrary multiplicity. The sum of all these partial cross sections equals that of the primary process. This property of the GPMCs reflects the KLN cancellation mentioned earlier, and it is often called “unitarity of the shower process”, a name that reminds us that the KLN cancellation itself is a consequence of unitarity. The fact that a quantum mechanical process can be described in terms of composition of probabilities, rather than amplitudes, follows from the LL approximation. In fact, in the dominant, strongly ordered region, subsequent splittings are separated by increasingly large times and distances, and this suppresses interference effects.

We now illustrate the basic parton-shower algorithm, as first introduced in Ref. 8. The purpose of this illustration is to give a schematic representation of how shower algorithms work, to introduce some concepts that will be referred to in the following, and to show the relationship between shower algorithms and Feynman-diagram results. For simplicity, we consider the example of  $e^+e^-$  annihilation into  $q\bar{q}$  pairs. With each dominant (i.e. strongly ordered) final-state configuration one can associate an ordered tree diagram, by recursively clustering together final-state parton pairs with the smallest hardness, and ending up with the hard production vertex (i.e. the  $\gamma^* \rightarrow q\bar{q}$ ). The momenta of all intermediate lines of the tree diagram are then uniquely determined from the final-state momenta. Hardnesses in the graph are also strongly ordered. One assigns to each splitting vertex the hardness  $t$ , the energy fractions  $z$  and  $1-z$  of the two generated partons, and the azimuth  $\phi$  of the splitting process with respect to the momentum of the incoming parton. For definiteness, we assume that  $z$  and  $\phi$  are defined in the center-of-mass (CM) frame of the  $e^+e^-$  collision, although other definitions are possible that differ only beyond the LL approximation. The differential cross section for a

given final state is given by the product of the differential cross section for the initial  $e^+e^- \rightarrow q\bar{q}$  process, multiplied by a factor

$$\Delta_i(t, t') \frac{\alpha_S(t)}{2\pi} P_{i,jk}(z) \frac{dt}{t} \frac{d\phi}{2\pi} \quad (41.3)$$

for each intermediate line ending in a splitting vertex. We have denoted with  $t'$  the maximal hardness that is allowed for the line, with  $t$  its hardness, and  $z$  and  $\phi$  refer to the splitting process.  $\Delta(t, t')$  is the so-called Sudakov form factor

$$\Delta_i(t, t') = \exp \left[ - \int_t^{t'} \frac{dq^2}{q^2} \frac{\alpha_S(q^2)}{2\pi} \sum_{jk} P_{i,jk}(z) dz \frac{d\phi}{2\pi} \right]. \quad (41.4)$$

The suffixes  $i$  and  $jk$  represent the parton species of the incoming and final partons, respectively, and  $P_{i,jk}(z)$  are the Altarelli-Parisi [9] splitting kernels. Final-state lines that do not undergo any further splitting are associated with a factor

$$\Delta_i(t_0, t'), \quad (41.5)$$

where  $t_0$  is an infrared cutoff defined by the shower hadronization scale (at which the charges are screened by hadronization) or, for an unstable particle, its width (a source cannot emit radiation with a period exceeding its lifetime).

Notice that the definition of the Sudakov form factor is such that

$$\Delta_i(t_2, t_1) + \int_{t_2}^{t_1} \frac{dt}{t} dz \frac{d\phi}{2\pi} \sum_{jk} \Delta_i(t, t_1) \frac{\alpha_S(t)}{2\pi} P_{i,jk}(z) = 1. \quad (41.6)$$

This implies that the cross section for developing the shower up to a given stage does not depend on what happens next, since subsequent factors for further splitting or not splitting add up to one.

The shower cross section can then be formulated in a probabilistic way. The Sudakov form factor  $\Delta_i(t_2, t_1)$  is interpreted as the probability for a splitting *not* to occur, for a parton of type  $i$ , starting from a branching vertex at the scale  $t_1$ , down to a scale  $t_2$ . Notice that  $0 < \Delta_i(t_2, t_1) \leq 1$ , where the upper extreme is reached for  $t_2 = t_1$ , and the lower extreme is approached for  $t_2 = t_0$ . From Eq. (41.4), it seems that the Sudakov form factor should vanish if  $t_2 = 0$ . However, because of the presence of the running coupling in the integrand,  $t_2$  cannot be taken smaller than some cutoff scale of the order of  $\Lambda$ , so that at its lower extreme the Sudakov form factor is small, but not zero. Event generation then proceeds as follows. One gets a uniform random number  $0 \leq r \leq 1$ , and seeks a solution of the equation  $r = \Delta_i(t_2, t_1)$  as a function of  $t_2$ . If  $r$  is too small and no solution exists, no splitting is generated, and the line is interpreted as a final parton. If a solution  $t_2$  exists, a branching is generated at the scale  $t_2$ . Its  $z$  value and the final parton species  $jk$  are generated with a probability proportional to  $P_{i,jk}(z)$ . The azimuth is generated uniformly. This procedure is started with each of the primary process partons, and is applied recursively to all generated partons. It may generate an arbitrary number of partons, and it stops when no final-state partons undergo further splitting.

The four-momenta of the final-state partons are reconstructed from the momenta of the initiating ones, and from the whole sequence of splitting variables, subject to overall momentum conservation. Different algorithms employ different strategies to treat recoil effects due to momentum conservation, which may be applied either locally for each parton or dipole splitting, or globally for the entire set of partons (a procedure called *momentum reshuffling*). This has a subleading effect with respect to the collinear approximation.

We emphasize that the shower cross section described above can be derived from perturbative QCD by keeping only the collinear-dominant real and virtual contributions to the cross section. In particular, up to terms that vanish after azimuthal averaging, the product of the cross section for the basic process, times the factors

$$\frac{\alpha_S}{2\pi} \frac{dt}{t} dz \frac{d\phi}{2\pi} P_{i,jk}(z) \quad (41.7)$$

at each branching vertex, gives the leading collinear contribution to the tree-level cross section for the same process. The dominant virtual corrections in the same approximation are provided by the running coupling at each vertex and by the Sudakov form factors in the intermediate lines.

#### 41.1.1. Angular correlations :

In gluon splitting processes ( $g \rightarrow q\bar{q}$ ,  $g \rightarrow gg$ ) in the collinear approximation, the distribution of the split pair is not uniform in azimuth, and the Altarelli-Parisi splitting functions are recovered only after azimuthal averaging. This dependence is due to the interference of positive and negative helicity states for the gluon that undergoes splitting. Spin correlations propagate through the splitting process, and determine acausal correlations of the EPR kind [10]. A method to partially account for these effects was introduced in Ref. 11, in which the azimuthal correlation between two successive splittings is computed by averaging over polarizations. This can then be applied at each branching step. Acausal correlations are argued to be small, and are discarded with this method, that is still used in the PYTHIA code [3]. A method that fully includes spin correlation effects was later proposed by Collins [12], and has been implemented in the fortran HERWIG code [13].

#### 41.1.2. Initial-state radiation :

Initial-state radiation (ISR) arises because incoming charged particles can radiate before entering the hard-scattering process. In doing so, they acquire a non-vanishing transverse momentum, and their virtuality becomes negative (spacelike). The dominant logarithmic region is the collinear one, where virtualities become larger and larger in absolute value with each emission, up to a limit given by the hardness of the basic process itself. A shower that starts by considering the highest virtualities first would thus have to work backward in time for ISR. A corresponding backwards-evolution algorithm was formulated by Sjöstrand [14], and was basically adopted in all shower models.

The key point in backwards evolution is that the evolution probability depends on the amount of partons that could have given rise to the one being evolved. This is reflected by introducing the ratio of the PDF after the branching to the PDF before the branching in the definition of the backward-evolution Sudakov form factor,

$$\Delta_i^{\text{ISR}}(t, t') = \exp \left[ - \int_{t'}^t \frac{dt''}{t''} \frac{\alpha_S(t'')}{2\pi} \int_x^1 \frac{dz}{z} \sum_{jk} P_{j,ik}(z) \frac{f_j(t'', x/z)}{f_i(t'', x)} \right]. \quad (41.8)$$

Notice that there are two uses of the PDFs: they are used to compute the cross section for the basic hard process, and they control ISR via backward evolution. Since the evolution is generated with leading-logarithmic accuracy, it is acceptable to use two different PDF sets for these two tasks, provided they agree at the LO level.

In the context of GPMC evolution, each ISR emission generates a finite amount of transverse momentum. Details on how the recoils generated by these transverse “kicks” are distributed among other partons in the event, in particular the ones involved in the hard process, constitute one of the main areas of difference between existing algorithms, see Ref. 15. An additional  $\mathcal{O}(1 \text{ GeV})$  of “primordial  $k_T$ ” is typically added, to represent the sum of unresolved and/or non-perturbative motion below the shower cutoff scale.

#### 41.1.3. Soft emissions and QCD coherence :

In massless field theories like QCD, there are two sources of large logarithms of infrared origin. One has to do with collinear singularities, which arise when two final-state particles become collinear, or when a final-state particle becomes collinear to an initial-state one. The other has to do with the emission of soft gluons at arbitrary angles. Because of that, it turns out that in QCD perturbation theory two powers of large logarithms can arise for each power of  $\alpha_S$ . The expansion in leading soft and collinear logarithms is often referred to as the double-logarithmic expansion.

Within the conventional parton-shower formalism, based on collinear factorization, it was shown in a sequel of publications (see Ref. 16 and references therein) that the double-logarithmic region can be correctly described by using the angle of the emissions as the ordering variable, rather than the virtuality, and that the argument of  $\alpha_S$  at the splitting vertex should be the relative parton transverse momentum after the splitting. Physically, the ordering in angle approximates the coherent interference arising from large-angle

soft emission from a bunch of collinear partons. Without this effect, the particle multiplicity would grow too rapidly with energy, in conflict with  $e^+e^-$  data. For this reason, angular ordering is used as the evolution variable in both the HERWIG [16] and HERWIG++ [17] programs, and an angular veto is imposed on the virtuality-ordered evolution in PYTHIA 6 [18].

A radical alternative formulation of QCD cascades first proposed in Ref. 19 focuses upon soft emission, rather than collinear emission, as the basic splitting mechanism. It then becomes natural to consider a branching process where it is a parton pair (i.e. a dipole) rather than a single parton, that emits a soft parton. Adding a suitable correction for non-soft, collinear partons, one can achieve in this framework the correct logarithmic structure for both soft and collinear emissions in the limit of large number of colors  $N_c$ , without any explicit angular-ordering requirement. The ARIADNE [20] and VINCIA [21] programs are based on this approach. In SHERPA, the default shower [22] is also of a dipole type [23], while the  $p_\perp$ -ordered showers in PYTHIA 6 and 8 represent a hybrid, combining collinear splitting kernels with dipole kinematics [24].

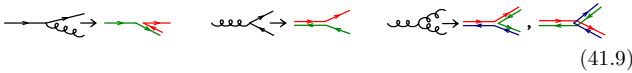
#### 41.1.4. Massive quarks :

Quark masses act as cut-off on collinear singularities. If the mass of a quark is below, or of the order of  $\Lambda$ , its effect in the shower is small. For larger quark masses, like in  $c$ ,  $b$ , or  $t$  production, it is the mass, rather than the typical hadronic scale, that cuts off collinear radiation. For a quark with energy  $E$  and mass  $m_Q$ , the divergent behavior  $d\theta/d\theta$  of the collinear splitting process is regulated for  $\theta \leq \theta_0 = m_Q/E$ . We thus expect less collinear activity for heavy quarks than for light ones, which in turn is the reason why heavy quarks carry a larger fraction of the momentum acquired in the hard production process.

This feature can be implemented with different levels of sophistication. Using the fact that soft emission exhibits a zero at zero emission angle, older parton shower algorithms simply limited the shower emission to be not smaller than the angle  $\theta_0$ . More modern approaches are used in both PYTHIA, where mass effects are included using a kind of matrix-element correction method [25], and in HERWIG++ and SHERPA, where a generalization of the Altarelli-Parisi splitting kernel is used for massive quarks [26].

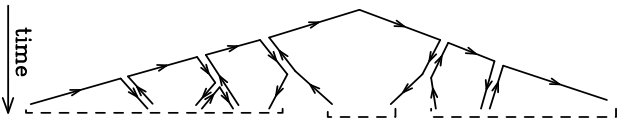
#### 41.1.5. Color information :

Shower MC generators track large- $N_c$  color information during the development of the shower. In the large- $N_c$  limit, quarks or antiquarks are represented by a color line, i.e. a line with an arrow indicating the direction of color flow. Gluons are represented by a pair of color lines with opposite arrows. The rules for color propagation are:



(41.9)

During the shower development, partons are connected by color lines. We can have a quark directly connected by a color line to an antiquark, or via an arbitrary number of intermediate gluons, as shown in Fig. 41.1.



**Figure 41.1:** Color development of a shower in  $e^+e^-$  annihilation. Systems of color-connected partons are indicated by the dashed lines.

It is also possible for a set of gluons to be connected cyclically in color, as e.g. in the decay  $\Upsilon \rightarrow ggg$ .

The color information is used in angular-ordered showers, where the angle of color-connected partons determines the initial angle for the shower development, and in dipole showers, where dipoles are always color-connected partons. It is also used in hadronization models, where the initial strings or clusters used for hadronization are formed by systems of color-connected partons.

#### 41.1.6. Electromagnetic corrections :

The physics of photon emission from light charged particles can also be treated with a shower MC algorithm. High-energy electrons and quarks, for example, are accompanied by bremsstrahlung photons. Also here, similarly to the QCD case, electromagnetic corrections are of order  $\alpha_{em} \ln(Q/m)$ , where  $m$  is the mass of the radiating particle, or even of order  $\alpha_{em} \ln(Q/m) \ln(E_\gamma/E)$  in the region where soft photon emission is important, so that, especially for the case of electrons, their inclusion in the simulation process is mandatory. This is done in most of the GPMC's (for a recent comparative study see [27]). The specialized generator PHOTOS [28] is sometimes used as an afterburner for an improved treatment of QED radiation in non-hadronic resonance decays.

In case of photons emitted by leptons the shower can be continued down to virtualities arbitrarily close to the lepton mass shell (unlike the case in QCD). In practice, photon radiation must be cut off below a certain energy, in order for the shower algorithm to terminate. Therefore, there is always a minimum energy for emitted photons that depends upon the implementations [27] (and so does the MC truth for a charged lepton). In the case of electrons, this energy is typically of the order of its mass. Electromagnetic radiation below this scale is not enhanced by collinear singularities, and is thus bound to be soft, so that the electron momentum is not affected by it.

For photons emitted from quarks, we have instead the obvious limitation that the photon wavelength cannot exceed the typical hadronic size. Longer-wavelength photons are in fact emitted by hadrons, rather than quarks. This last effect is in practice never modeled by existing shower MC implementations. Thus, electromagnetic radiation from quarks is cut off at a typical hadronic scale. Finally, hadron (and  $\tau$ ) decays involving charged particles can produce additional soft bremsstrahlung. This is implemented in a general way in HERWIG++ [29] and SHERPA [30].

#### 41.1.7. Beyond-the-Standard-Model Physics :

The inclusion of processes for physics beyond the Standard Model (BSM) in event generators is to some extent only a matter of implementing the relevant hard processes and (chains of) decays, with the level of difficulty depending on the complexity of the model and the degree of automation [31,32]. Notable exceptions are long-lived colored particles [33], particles in exotic color representations, and particles showering under new gauge symmetries, with a growing set of implementations documented in the individual GPMC manuals. Further complications that may be relevant are finite-width effects (discussed in Sec. 41.1.8) and the assumed threshold behavior.

In addition to code-specific implementations [15], there are a few commonly adopted standards that are useful for transferring information and events between codes. Currently, the most important of these is the Les Houches Event File (LHEF) standard [34], normally used to transfer parton-level events from a hard-process generator to a shower generator. Another important standard is the Supersymmetry Les Houches Accord (SLHA) format [35], originally used to transfer information on supersymmetric particle spectra and couplings, but by now extended to apply also to more general BSM frameworks and incorporated within the LHEF standard [36].

#### 41.1.8. Decay Chains and Particle Widths :

In most BSM processes and some SM ones, an important aspect of the event simulation is how decays of short-lived particles, such as top quarks, EW and Higgs bosons, and new BSM resonances, are handled. We here briefly summarize the spectrum of possibilities, but emphasize that there is no universal standard. Users are advised to check whether the treatment of a given code is adequate for the physics study at hand.

The appearance of an unstable resonance as a physical particle at some intermediate stage of the event generation implies that its production and decay processes are treated as being factorized. This is valid up to corrections of order  $\Gamma/m_0$ , with  $\Gamma$  the width and  $m_0$  the pole mass. States whose widths are a substantial fraction of their mass should not be treated as “physical particles,” but rather as intrinsically off-shell internal propagator lines.

For states treated as physical particles, two aspects are relevant: the mass distribution of the decaying particle itself and the distributions

of its decay products. For the former, matrix-element generators often use a simple  $\delta$  function at  $m_0$ . The next level up, typically used in GPMCs, is to use a Breit-Wigner distribution (relativistic or non-relativistic), which formally resums higher-order virtual corrections to the mass distribution. Note, however, that this still only generates an improved picture for *moderate* fluctuations away from  $m_0$ . Similarly to above, particles that are significantly off-shell (in units of  $\Gamma$ ) should not be treated as resonant, but rather as internal off-shell propagator lines. In most GPMCs, further refinements are included, for instance by letting  $\Gamma$  be a function of  $m$  (“running widths”) and by limiting the magnitude of the allowed fluctuations away from  $m_0$ .

For the distributions of the decay products, the simplest treatment is again to assign them their respective  $m_0$  values, with a uniform phase-space distribution. A more sophisticated treatment distributes the decay products according to the differential decay matrix elements, capturing at least the internal dynamics and helicity structure of the decay process, including EPR-like correlations. Further refinements include polarizations of the external states [37] and assigning the decay products their own Breit-Wigner distributions, the latter of which opens the possibility to include also intrinsically off-shell decay channels, like  $H \rightarrow WW^*$ .

During subsequent showering of the decay products, most parton-shower models will preserve their total invariant mass, so as not to skew the original resonance shape.

When computing partial widths and/or modifying decay tables, one should be aware of the danger of double-counting intermediate on-shell particles, see Sec. 41.2.3.

#### 41.1.9. Matching with Matrix Elements :

Shower algorithms are based upon a combination of the collinear (small-angle) and soft (small-energy) approximations and are thus normally inaccurate for hard, wide-angle emissions (i.e., additional well-resolved jets). They also contain only the leading singular pieces of next-to-leading order (NLO) and higher corrections to the basic process.

Traditional GPMCs, like HERWIG and PYTHIA, have included for a long time the so called Matrix Element Corrections (MEC), first formulated in Ref. 38 with later developments summarized in Ref. 15. They are typically available for  $2 \rightarrow 1$  or  $1 \rightarrow 2$  processes, like DIS, vector boson and Higgs production and decays, and top decays. The MEC corrects the emission of the hardest jet at large angles, so that it becomes exact at LO. A generalization of the method to multiple emissions was formulated recently [39].

Aside from MECs implemented directly in the GPMCs, the improvements on the parton-shower description of hard collisions have been made in two main directions: the so called Matrix Elements and Parton Shower matching (ME+PS from now on), and the matching of NLO calculations and Parton Showers (NLO+PS). We now discuss each of these, and then briefly summarise techniques becoming available for combining them.

The ME+PS method allows one to use tree-level matrix elements for hard, large-angle emissions. It was first formulated in the so-called CKKW paper [40], and several variants have appeared, including the CKKW-L, MLM, and pseudoshower methods, see Refs. 41, 15 for summaries. Truncated showers are required [42] to maintain color coherence when interfacing to angular-ordered parton showers, and care must be taken to use consistent  $\alpha_s$  choices for the real (ME-driven) and virtual (PS-driven) corrections [43].

In the ME+PS method one typically starts by generating LO matrix elements for the production of the basic process plus a certain number  $\leq n$  of other partons. A minimum separation is imposed on the produced partons, requiring, for example, that the relative transverse momentum in any pair of partons is above a given cut  $Q_{\text{cut}}$ . One then reweights these amplitudes in such a way that, in the strongly ordered region, the virtual effects that are included in the shower algorithm (i.e. running couplings and Sudakov form factors) are also accounted for. At this stage, before parton showers are added, the generated configurations are tree-level accurate at large angle, and at small angle they match the results of the shower algorithm, except that there are no emissions below the scale  $Q_{\text{cut}}$ , and no final

states with more than  $n$  partons. These kinematic configurations are thus fed into a GPMC, that must generate all splittings with relative transverse momentum below the scale  $Q_{\text{cut}}$ , for initial events with less than  $n$  partons, or below the scale of the smallest pair transverse momentum, for events with  $n$  partons. The matching parameter  $Q_{\text{cut}}$  must be chosen to be large enough for fixed-order perturbation theory to hold, but small enough so that the shower is accurate for emissions below it. Notice that the accuracy achieved with MEC is equivalent to that of ME+PS with  $n = 1$ , where MEC has the advantage of not having a matching parameter  $Q_{\text{cut}}$ .

The popularity of the ME+PS method is due to the fact that processes with many jets appear often as backgrounds to new-physics searches. These jets are typically required to be well separated, and to have large transverse momenta. These kinematical configurations are exactly those for which pure shower algorithms are unreliable, hence it is mandatory to describe them using at least LO matrix elements.

Several ME+PS implementations use existing LO generators, like ALPGEN [52], MADGRAPH [53], and others summarized in Ref. 41, for the calculation of the matrix elements, and feed the partonic events to a GPMC like PYTHIA or HERWIG using the Les Houches Interface for User Processes (LHI/LHEF) [54,34]. SHERPA and HERWIG++ also include their own matrix-element generators.

The NLO+PS methods promote the accuracy of the generation of the basic process from LO to NLO in QCD. They must thus include the radiation of one extra parton with tree-level accuracy, since this radiation constitutes a NLO correction to the basic process. They must also include NLO virtual corrections. They can be viewed as an extension of the MEC methods with the inclusion of NLO virtual corrections. They are however more general, since they are applicable to processes of arbitrary complexity. Two of these methods are now widely used: MC@NLO [44] and POWHEG [42,45], with several alternative methods now also being pursued, see Ref. 15 and references therein.

NLO+PS generators produce NLO accurate distributions for inclusive quantities, and generate the hardest jet with tree-level accuracy. It should be recalled, though, that in  $2 \rightarrow 1$  processes like  $Z/W$  production, GPMCs including MEC and weighted by a constant  $K$  factor may perform nearly as well, and, if suitably tuned, may even yield a better description of data. In this context, note also that the optimal tuning of an NLO+PS generator may well be different from that of the pure PS.

Several NLO+PS processes are implemented in the MC@NLO program [44], together with the new AMC@NLO development [55], and in the POWHEG BOX framework [45]. HERWIG++ also includes its own POWHEG implementation, suitably adapted with the inclusion of vetoed and truncated showers, for several processes. SHERPA instead implements a variant of the MC@NLO method.

For applications that require an accurate description of more than one hard, large-angle jet associated with the primary process, ME+PS schemes are still superior to NLO+PS ones. Ideally, one would like to improve NLO generators in such a way that also the production of associated jets achieves NLO accuracy. The FxFX [47], UNLOPS [48], MiNLO [49] and MEPS@NLO [50] methods address this problem. In turn, its solution is a prerequisite for the construction of NNLO+PS generators, that in fact have already appeared for the  $gg \rightarrow H$  and Drell-Yan processes (see ref. [51] and references therein).

## 41.2. Hadronization Models

In the context of GPMCs, *hadronization* denotes the process by which a set of colored partons (*after* showering) is transformed into a set of color-singlet *primary* hadrons, which may then subsequently decay further (to *secondary* hadrons). This non-perturbative transition takes place at the *hadronization scale*  $Q_{\text{had}}$ , which by construction is identical to the infrared cutoff of the parton shower. In the absence of a first-principles solution to the relevant dynamics, GPMCs use QCD-inspired phenomenological models to describe this transition.

A key difference between MC hadronization models and the fragmentation-function (FF) formalism used to describe inclusive hadron spectra in perturbative QCD (see Chap. 9 and Chap. 20 of



PDG book) is that the former is always defined at the hadronization scale, while the latter can be defined at an arbitrary perturbative scale  $Q$ . They can therefore only be compared directly if the perturbative evolution between  $Q$  and  $Q_{\text{had}}$  is taken into account. FFs are calculable in pQCD, given a non-perturbative initial condition obtained by fits to hadron spectra. In the MC context, one can prove that the correct QCD evolution of the FFs arises from the shower formalism, with the hadronization model providing an explicit parametrization of the non-perturbative component. However, the MC modeling of shower and hadronization includes much more information on the final state since it is fully exclusive (i.e., it addresses all particles in the final state explicitly), while FFs only describe inclusive spectra. This exclusivity also enables MC models to make use of the color-flow information coming from the perturbative shower evolution (see Sec. 41.1.5) to determine between which partons the confining potentials should arise. This is the non-perturbative analog of QCD coherence [56].

Given an exact hadronization model, its dependence on the hadronization scale  $Q_{\text{had}}$  should in principle be compensated by the corresponding scale dependence of the shower algorithm, which stops generating branchings at the scale  $Q_{\text{had}}$ . However, due to their complicated and fully exclusive nature, it is generally not possible to enforce this compensation automatically in MC models. One must therefore be aware that the model must be “retuned” by hand if changes are made to the perturbative evolution, in particular if the infrared cutoff is modified. Tuning is discussed briefly in Sec. 41.4.

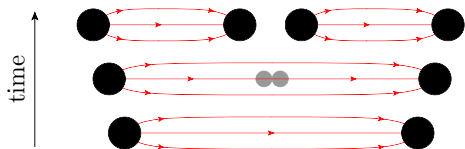
An important result in “quenched” lattice QCD (see Chap. 18 of PDG book) is that the potential of the color-dipole field between a charge and an anticharge appears to grow linearly with the separation of the charges, at distances greater than about a femtometer. This is known as “linear confinement”, and it forms the starting point for the *string model of hadronization*, discussed below in Sec. 41.2.1. Alternatively, a property of perturbative QCD called “preconfinement” is the basis of the *cluster model of hadronization*, discussed in Sec. 41.2.2.

Finally, it should be emphasized that the so-called “parton level” that can be obtained by switching off hadronization in a GPMC, is not a universal concept, since each model defines the hadronization scale differently (e.g. by a cutoff in  $p_{\perp}$ , invariant mass, etc., with different tunes using different values for the cutoff). Comparisons to distributions at this level may therefore be used to provide an idea of the overall impact of hadronization corrections within a given model, but should be avoided in the context of physical observables.

#### 41.2.1. The String Model :

Starting from early concepts [57], several hadronization models based on strings have been proposed [15]. Of these, the most widely used today is the so-called Lund model [58,59], implemented in PYTHIA [3,4]. We concentrate on that particular model here, though many of the overall concepts would be shared by any string-inspired method.

Consider a color-connected quark-antiquark pair with no intermediate gluons emerging from the parton shower (like the  $q\bar{q}$  pair in the center of Fig. 41.1), e.g. a red  $q$  and an antired  $\bar{q}$ . As the charges move apart, linear confinement implies that a potential  $V(r) = \kappa r$  is reached for large distances  $r$ . (At short distances, there is a Coulomb term  $\propto 1/r$  as well, but this is neglected in the Lund string.) This potential describes a string with tension  $\kappa \sim 1 \text{ GeV/fm} \sim 0.2 \text{ GeV}^2$ . The physical picture is that of a color flux tube being stretched between the  $q$  and the  $\bar{q}$ .



**Figure 41.2:** Illustration of string breaking by quark pair-creation in the string field.

As the string grows, the non-perturbative creation of quark-antiquark pairs can break the string, via the process  $(q\bar{q}) \rightarrow (q\bar{q}') + (q'\bar{q})$ , illustrated in Fig. 41.2. More complicated color-connected quark-antiquark configurations involving intermediate gluons (like the  $\bar{q}gggq$  and  $\bar{q}gq$  systems on the left and right part of Fig. 41.1) are treated by representing gluons as transverse “kinks.” Thus soft gluons effectively build up a transverse structure in the originally one-dimensional object, with infinitely soft ones smoothly absorbed into the string. For strings with finite-energy kinks, the space-time evolution is slightly more involved [59], but the main point is that there are no separate free parameters for gluon jets. Differences with respect to quark fragmentation arise simply because quarks are only connected to a single string piece, while gluons have one on either side, increasing their relative energy loss (per unit invariant time) by a factor of 2, similar to the ratio of color Casimirs  $C_A/C_F = 2.25$ .

Since the string breaks are causally disconnected (as can be realized from space-time diagrams [59]), they do not have to be considered in any specific time-ordered sequence. In the Lund model, the string breaks are generated starting with the leading (“outermost”) hadrons, containing the endpoint quarks, and iterating inwards towards the center of the string, alternating randomly between the left and right sides. One can thereby split off a single on-shell hadron in each step, making it straightforward to ensure that only states consistent with known hadron states are produced.

For each breakup vertex, quantum mechanical tunneling is assumed to control the masses and  $p_{\perp}$  kicks that can be produced, leading to a Gaussian suppression

$$\text{Prob}(m_q^2, p_{\perp q}^2) \propto \exp\left(\frac{-\pi m_q^2}{\kappa}\right) \exp\left(\frac{-\pi p_{\perp q}^2}{\kappa}\right), \quad (41.10)$$

where  $m_q$  is the mass of the produced quark flavor and  $p_{\perp}$  is the non-perturbative transverse momentum imparted to it by the breakup process (the antiquark has the same mass and opposite  $p_{\perp}$ ), with a universal average value of  $\langle p_{\perp q}^2 \rangle = \kappa/\pi \sim (250 \text{ MeV})^2$ . The charm and bottom masses are sufficiently heavy that they are not produced at all in the soft fragmentation. The transverse direction is defined with respect to the string axis, so the  $p_{\perp}$  in a frame where the string is moving will be modified by a Lorentz boost. Note that the effective amount of “non-perturbative”  $p_{\perp}$ , in a Monte Carlo model with a fixed shower cutoff  $Q_{\text{had}}$ , may be larger than the purely non-perturbative  $\kappa/\pi$  above, to account for effects of additional unresolved soft-gluon radiation below  $Q_{\text{had}}$ . In principle, the magnitude of this additional component should scale with the cutoff, but in practice it is up to the user to enforce this by retuning the relevant parameter when changing the hadronization scale.

Since quark masses are difficult to define for light quarks, the value of the strangeness suppression is determined from experimental observables, such as the  $K/\pi$  and  $K^*/\rho$  ratios. Note that the parton-shower evolution generates a small amount of strangeness as well, through perturbative  $g \rightarrow s\bar{s}$  splittings.

Baryon production can also be incorporated, by allowing string breaks to produce pairs of *diquarks*, loosely bound states of two quarks in an overall  $\bar{3}$  representation. Again, since diquark masses are difficult to define, the relative rate of diquark to quark production is extracted, e.g. from the  $p/\pi$  ratio. Since the perturbative shower splittings do not produce diquarks, the optimal value for this parameter is mildly correlated with the amount of  $g \rightarrow q\bar{q}$  splittings produced by the shower. More advanced scenarios for baryon production have also been proposed, see Ref. 59. Within the PYTHIA framework, a fragmentation model including baryon string junctions [60] is also available.

The next step of the algorithm is the assignment of the produced quarks within hadron multiplets. Using a nonrelativistic classification of spin states, the fragmenting  $q$  may combine with the  $\bar{q}'$  from a newly created breakup to produce a meson — or baryon, if diquarks are involved — of a given spin  $S$  and angular momentum  $L$ . The lowest-lying pseudoscalar and vector meson multiplets, and spin-1/2 and -3/2 baryons, are assumed to dominate in a string framework<sup>1</sup>,

<sup>1</sup> The PYTHIA implementation includes the lightest pseudoscalar and



but individual rates are not predicted by the model. This is therefore the sector that contains the largest amount of free parameters.

From spin counting, the ratio  $V/P$  of vectors to pseudoscalars is expected to be 3, but in practice this is only approximately true for  $B$  mesons. For lighter flavors, the difference in phase space caused by the  $V$ - $P$  mass splittings implies a suppression of vector production. When extracting the corresponding parameters from data, it is advisable to begin with the heaviest states, since so-called feed-down from the decays of higher-lying hadron states complicates the extraction for lighter particles, see Sec. 41.2.3. For baryons, separate parameters control the relative rates of spin-1 diquarks vs. spin-0 ones and, likewise, have to be extracted from data.

With  $p_{\perp}^2$  and  $m^2$  now fixed, the final step is to select the fraction,  $z$ , of the fragmenting endpoint quark's longitudinal momentum that is carried by the created hadron, an aspect for which the string model is highly predictive. The requirement that the fragmentation be independent of the sequence in which breakups are considered (causality) imposes a "left-right symmetry" on the possible form of the fragmentation function,  $f(z)$ , with the solution

$$f(z) \propto \frac{1}{z}(1-z)^a \exp\left(-\frac{b(m_h^2 + p_{\perp h}^2)}{z}\right), \quad (41.11)$$

which is known as the Lund symmetric fragmentation function (normalized to unit integral). The dimensionless parameter  $a$  dampens the hard tail of the fragmentation function, towards  $z \rightarrow 1$ , and may in principle be flavor-dependent, while  $b$ , with dimension  $\text{GeV}^{-2}$ , is a universal constant related to the string tension [59] which determines the behavior in the soft limit,  $z \rightarrow 0$ . Note that the explicit mass dependence in  $f(z)$  implies a harder fragmentation function for heavier hadrons (in the rest frame of the string).

As a by-product, the probability distribution in invariant time  $\tau$  of  $q'\bar{q}$  breakup vertices, or equivalently  $\Gamma = (\kappa\tau)^2$ , is also obtained, with  $dP/d\Gamma \propto \Gamma^a \exp(-b\Gamma)$  implying an area law for the color flux, and the average breakup time lying along a hyperbola of constant invariant time  $\tau_0 \sim 10^{-23}\text{s}$  [59].

For massive endpoints (e.g.  $c$  and  $b$  quarks, or hypothetical hadronizing new-physics particles), which do not move along straight lightcone sections, the exponential suppression with string area leads to modifications of the form  $f(z) \rightarrow f(z)/z^{bm_Q^2}$ , with  $m_Q$  the mass of the heavy quark [61]. Although different forms can also be used to describe inclusive heavy-meson spectra (see Sec 20.9 of PDG book), such choices are not consistent with causality in the string framework and hence are theoretically disfavored in this context, one well-known example being the Peterson formula [62],

$$f(z) \propto \frac{1}{z} \left(1 - \frac{1}{z} - \frac{\epsilon_Q}{1-z}\right)^{-2}, \quad (41.12)$$

with  $\epsilon_Q$  a free parameter expected to scale  $\propto 1/m_Q^2$ .

#### 41.2.2. The Cluster Model:

The cluster hadronization model is based on *preconfinement*, i.e., on the observation [63,64] that the color structure of a perturbative QCD shower evolution at any scale  $Q_0$  is such that color-singlet subsystems of partons (labeled "clusters") occur with a universal invariant mass distribution that only depends on  $Q_0$  and on  $\Lambda_{\text{QCD}}$ , not on the starting scale  $Q$ , for  $Q \gg Q_0 \gg \Lambda_{\text{QCD}}$ . Further, this mass distribution is power-suppressed at large masses.

Following early models based on this universality [8,65], the cluster model developed by Webber [66] has for many years been a hallmark of the HERWIG and HERWIG++ generators, with an alternative implementation [67] now available in the SHERPA generator. The key idea, in addition to preconfinement, is to force "by hand" all gluons

vector mesons, with the four  $L = 1$  multiplets (scalar, tensor, and 2 pseudovectors) available but disabled by default, largely because several states are poorly known and thus may result in a worse overall description when included. For baryons, the lightest spin-1/2 and -3/2 multiplets are included.

to split into quark-antiquark pairs at the end of the parton shower. Compared with the string description, this effectively amounts to viewing gluons as "seeds" for string breaks, rather than as kinks in a continuous object. After the splittings, a new set of low-mass color-singlet clusters is obtained, formed only by quark-antiquark pairs. These can be decayed to on-shell hadrons in a simple manner.

The algorithm starts by generating the forced  $g \rightarrow q\bar{q}$  breakups, and by assigning flavors and momenta to the produced quark pairs. For a typical shower cutoff corresponding to a gluon virtuality of  $Q_{\text{had}} \sim 1\text{GeV}$ , the  $p_{\perp}$  generated by the splittings can be neglected. The constituent light-quark masses,  $m_{u,d} \sim 300\text{MeV}$  and  $m_s \sim 450\text{MeV}$ , imply a suppression (typically even an absence) of strangeness production. In principle, the model also allows for diquarks to be produced at this stage, but due to the larger constituent masses this would only become relevant for shower cutoffs larger than  $1\text{GeV}$ .

If a cluster formed in this way has an invariant mass above some cutoff value, typically 3–4 GeV, it is forced to undergo sequential  $1 \rightarrow 2$  cluster breakups, along an axis defined by the constituent partons of the original cluster, until all sub-cluster masses fall below the cutoff value. Due to the preservation of the original axis in these breakups, this treatment has some resemblance to the string-like picture, though the non-perturbative  $p_{\perp}$  kicks generated in this way are generally larger, up to half the allowed cluster mass.

Next, on the low-mass side of the spectrum, some clusters are allowed to decay directly to a single hadron, with nearby clusters absorbing any excess momentum. This improves the description of the high- $z$  part of the fragmentation spectrum — where the hadron carries almost all the momentum of its parent jet — at the cost of introducing one additional parameter, controlling the probability for single-hadron cluster decay.

Having obtained a final distribution of small-mass clusters, now with a strict cutoff at 3–4 GeV and with the component destined to decay to single hadrons already removed, the remaining clusters are interpreted as a smoothed-out spectrum of excited mesons, each of which decays isotropically to two hadrons, with relative probabilities proportional to the available phase space for each possible two-hadron combination that is consistent with the cluster's internal flavors, including spin degeneracy. It is important that all the light members (containing only  $uds$ ) of each hadron multiplet be included, as the absence of members can lead to unphysical isospin or  $\text{SU}(3)$  flavor violation. Typically, the lightest pseudoscalar, vector, scalar, even and odd charge conjugation pseudovector, and tensor multiplets of light mesons are included. In addition, some excited vector multiplets of light mesons may be available. For baryons, usually only the lightest flavor-octet, -decuplet and -singlet baryons are present, although both the HERWIG++ and SHERPA implementations now include some heavier baryon multiplets as well.

Differently from the string model, the mechanism of phase-space suppression employed here leads to a natural enhancement of the lighter pseudoscalars, and no parameters beyond the spectrum of hadron masses need to be introduced at this point. The phase space also limits the transverse momenta of the produced hadrons relative to the jet axis.

Note that, since the masses and decays of excited heavy-flavor hadrons in particular are not well known, there is some freedom in the model to adjust these, which in turn will affect their relative phase-space populations.

#### 41.2.3. Hadron and $\tau$ Decays:

Of the so-called primary hadrons, originating directly from string breaks and/or cluster decays (see above), many are unstable and so decay further, until a set of particles is obtained that can be considered stable on time scales relevant to the given measurement<sup>2</sup>. The decay modeling can therefore have a significant impact on final particle yields and spectra, especially for the lowest-lying hadronic

<sup>2</sup> E.g., a typical hadron-collider definition of a "stable particle" is  $c\tau \geq 10\text{mm}$ , which includes the weakly-decaying strange hadrons ( $K$ ,  $\Lambda$ ,  $\Sigma^{\pm}$ ,  $\bar{\Sigma}^{\pm}$ ,  $\Xi$ ,  $\Omega$ ).

states, which receive the largest relative contributions from decays (feed-down). Note that the interplay between primary production and feed-down implies that the hadronization parameters should be returned if significant changes to the decay treatment are made.

Particle summary tables, such as those given elsewhere in this *Review*, represent a condensed summary of the available experimental measurements and hence may be incomplete and/or exhibit inconsistencies within the experimental precision. In an MC decay package, on the other hand, all information must be quantified and consistent, with all branching ratios summing to unity. When adapting particle summary information for use in a decay package, a number of choices must therefore be made. The amount of ambiguity increases as more excited hadron multiplets are added to the simulation, about which less and less is known from experiment, with each GPMC making its own choices.

A related choice is how to distribute the decay products differentially in phase space, in particular which matrix elements to use. Historically, MC generators contained matrix elements only for selected (generator-specific) classes of hadron and  $\tau$  decays, coupled with a Breit-Wigner smearing of the masses, truncated at the edges of the physical decay phase space (the treatment of decay thresholds can be important for certain modes [15]). A more sophisticated treatment can then be obtained by reweighting the generated events using the obtained particle four-momenta and/or by using specialized external packages such as EVTGEN [68] for hadron decays and TAUOLA [69] for  $\tau$  decays.

More recently, HERWIG++ and SHERPA include helicity-dependence in  $\tau$  decays [70,71], with a more limited treatment available in PYTHIA 8 [4]. The HERWIG++ and SHERPA generators have also included significantly improved internal simulations of hadronic decays, which include spin correlations between those decays for which matrix elements are used. Photon-bremsstrahlung effects are discussed in Sec. 41.1.6.

HERWIG++ and PYTHIA include the probability for  $B$  mesons to oscillate into  $\bar{B}$  ones before decay. SHERPA and EVTGEN also include CP-violating effects and, for common decay modes of the neutral meson and its antiparticle, the interference between the direct decay and oscillation followed by decay.

We end on a note of warning on double counting. This may occur if a particle can decay via an intermediate on-shell resonance. An example is  $a_1 \rightarrow \pi\pi\pi$  which may proceed via  $a_1 \rightarrow \rho\pi$ ,  $\rho \rightarrow \pi\pi$ . If these decay channels of the  $a_1$  are both included, each with their full partial width, a double counting of the on-shell  $a_1 \rightarrow \rho\pi$  contribution would result. Such cases are normally dealt with consistently in the default MC generator packages, so this warning is mostly for users that wish to edit decay tables on their own.

## 41.3. Models for Soft Hadron-Hadron Physics

### 41.3.1. Minimum-Bias and Diffraction :

The term “minimum bias” (MB) originates from the experimental requirement of a minimal number of tracks (or hits) in a given instrumented region. In order to make MC predictions for such observables, all possible contributions to the relevant phase-space region must be accounted for. There are essentially four types of physics processes, which together make up the total hadron-hadron ( $hh$ ) cross section: 1) elastic scattering<sup>3</sup>:  $hh \rightarrow hh$ , 2) single diffractive dissociation:  $hh \rightarrow h + \text{gap} + X$ , with  $X$  denoting anything that is not the original beam particle, and “gap” denoting a rapidity region devoid of observed activity; 3) double diffractive dissociation:  $hh \rightarrow X + \text{gap} + X$ , and 4) inelastic non-diffractive scattering: everything else. A fifth class may also be defined, called central diffraction ( $hh \rightarrow h + \text{gap} + X + \text{gap} + h$ ). Some differences exist between theoretical and experimental terminology [72]. In the experimental setting, diffraction is defined by an observable gap, of some minimal size in rapidity. In the MC context, each diffractive physics process typically produces a whole spectrum of gaps, with small ones suppressed but not excluded.

<sup>3</sup> The QED elastic-scattering cross section diverges and is normally a non-default option in MC models.

The inelastic non-diffractive part of the cross section is typically modeled either by smoothly regulating and extending the perturbative QCD scattering cross sections all the way to zero  $p_\perp$  [73] (PYTHIA 6, PYTHIA 8, and SHERPA), or by regulating the QCD cross sections with a sharp cutoff [74] (HERWIG+JIMMY) and adding a separate class of intrinsically soft scatterings below that scale [75] (HERWIG++). See also Sec. 41.3.2. In all cases, the three most important ingredients are: 1) the IR regularization of the perturbative scattering cross sections, including their PDF dependence, 2) the assumed matter distribution of the colliding hadrons, possibly including multi-parton correlations [60] and/or  $x$  dependence [76], and 3) additional soft-QCD effects such as color reconnections and/or other collective effects, discussed in Sec. 41.3.3.

Currently, there are essentially three methods for simulating diffraction in the main MC models: 1) in PYTHIA 6, one picks a diffractive mass according to parametrized cross sections  $\propto dM^2/M^2$  [77]. This mass is represented as a string, which is fragmented as described in Sec. 41.2.1, though differences in the effective scale of the hadronization may necessitate a (re)tuning of the fragmentation parameters for diffraction; 2) in PYTHIA 8, the high-mass tail beyond  $M \sim 10$  GeV is augmented by a partonic description in terms of pomeron PDFs [78], allowing diffractive jet production including showers and underlying event [79]; 3) the PHOJET and DPMJET programs also include central diffraction and rely directly on a formulation in terms of pomerons (color-singlet multi-gluon states) [80–82]. Cut pomerons correspond to exchanges of soft gluons while uncut ones give elastic and diffractive topologies as well as virtual corrections that help preserve unitarity. So-called “hard pomerons” provide a transition to the perturbative regime. Fragmentation is still handled using the Lund string model, so there is some overlap with the above models at the hadronization stage. In addition, a pomeron-based package exists for HERWIG [83], and an effort is underway to construct an MC implementation of the “KMR” model [84] within the SHERPA generator. Color reconnections (Sec. 41.3.3) may also play a role in creating rapidity gaps and the underlying event (Sec. 41.3.2) in destroying them.

### 41.3.2. Underlying Event and Jet Pedestals :

In the GPMC context, “underlying event” (UE) denotes any additional activity *beyond* the basic process and its associated ISR and FSR activity. The dominant contribution is believed to come from additional color exchanges between the beam particles, which in models are represented either as multiple parton-parton interactions (MPI) or as so-called cut pomerons (Sec. 41.3.1). The experimentally observed fact that the UE is more active than average (MB) events at the same CM energy is called the “jet pedestal” effect.

The most clearly identifiable consequence of MPI is arguably the possibility of observing several hard parton-parton interactions in one and the same hadron-hadron event. In the most likely case that they are all QCD  $2 \rightarrow 2$  interactions, this produces two or more back-to-back jet pairs, with each pair having a small value of  $\sum(\vec{p}_\perp)$ . The fraction of MPI that give rise to additional reconstructible jets is, however, small. Soft interactions, that exchange color and a small amount of momentum without giving rise to observable jets, are much more plentiful, and can give significant corrections to the color flow and total scattered energy of the event. This affects the final-state activity in a more global way, increasing hadron-multiplicity and summed  $E_T$  distributions, and contributing to the break-up of the beam remnants in the forward direction.

The first detailed Monte Carlo model for perturbative MPI was proposed in Ref. 73, and with some variation this still forms the basis for most modern implementations. Some useful additional references can be found in Ref. 15. The first crucial observation is that the  $t$ -channel propagators appearing in perturbative QCD  $2 \rightarrow 2$  scattering almost go on shell at low  $p_\perp$ , causing the differential cross sections to behave roughly as

$$d\sigma_{2 \rightarrow 2} \propto \frac{dt}{t^2} \sim \frac{dp_\perp^2}{p_\perp^4} . \quad (41.13)$$

This cross section represents the inclusive scattering of partons against partons in perturbative QCD, summed over all partons. Thus, if a single hadron-hadron scattering contains *two* parton-parton

interactions, that event will contribute twice to the parton-parton cross section  $\sigma_{2 \rightarrow 2}$  but only once to the hadron-hadron one  $\sigma_{\text{tot}}$ , and so on. In the limit that all the parton-parton interactions are independent and equivalent, one has

$$\sigma_{2 \rightarrow 2} = \langle n \rangle \sigma_{\text{tot}}, \quad (41.14)$$

with  $\langle n \rangle$  the average number of parton-parton interactions, typically defined with some minimal  $p_{\perp} > p_{\perp \text{min}}$  to render the parton-parton cross section finite. The probability for having  $n$  parton-parton scatterings in a single hadron-hadron collision then follows a Poisson distribution,

$$\mathcal{P}_n = \langle n \rangle^n \frac{\exp(-\langle n \rangle)}{n!}. \quad (41.15)$$

This simple argument expresses unitarity; instead of the total hadron-hadron interaction cross section diverging as the parton-parton  $p_{\perp} \rightarrow 0$  (which would violate unitarity), we have restated the problem so that it is now the *number of parton-parton interactions per hadron-hadron collision* that diverges, with the total hadron-hadron cross section remaining finite. At LHC energies, the parton-parton scattering cross sections computed using the LO QCD cross section folded with modern PDFs become larger than the total  $pp$  one for  $p_{\perp \text{min}}$  values of order 4–5 GeV (see e.g. [85,86]). One therefore expects the average number of perturbative MPI to exceed unity at around that scale.

Two important ingredients remain to fully regulate the remaining divergence. Firstly, the interactions cannot use up more momentum than is available in the parent hadron. This suppresses the large- $n$  tail of the estimate above. In PYTHIA-based models, the MPI are ordered in  $p_{\perp}$ , and the parton densities for each successive interaction are explicitly constructed so that the sum of  $x$  fractions can never be greater than unity. In the HERWIG models, the Poisson estimate of  $\langle n \rangle$  above is used as an initial guess, but the generation of actual MPI is stopped once the energy-momentum conservation limit is reached. Both of these approaches generate momentum (conservation) correlations among the MPI.

The second ingredient invoked to suppress the number of interactions, at low  $p_{\perp}$  and  $x$ , is color screening; if the wavelength  $\sim 1/p_{\perp}$  of an exchanged colored parton becomes larger than a typical color-anticolor separation distance, it will only see an *average* color charge that vanishes in the limit  $p_{\perp} \rightarrow 0$ , hence leading to suppressed interactions. This provides an infrared cutoff for MPI similar to that provided by the hadronization scale for parton showers. A first estimate of the color-screening cutoff would be the proton size,  $p_{\perp \text{min}} \approx \hbar/r_p \approx 0.3 \text{ GeV} \approx \Lambda_{\text{QCD}}$ , but empirically this appears to be far too low. In current models, one replaces the proton radius  $r_p$  in the above formula by a “typical color screening distance,” i.e., an average size of a region within which the net compensation of a given color charge occurs. This number is not known from first principles [84] and is perceived of simply as an effective cutoff parameter. The simplest choice is to introduce a step function  $\Theta(p_{\perp} - p_{\perp \text{min}})$ . Alternatively, one may note that the jet cross section is divergent like  $\alpha_s^2(p_{\perp}^2)/p_{\perp}^4$ , cf. Eq. (41.13), and that therefore a factor

$$\frac{\alpha_s^2(p_{\perp 0}^2 + p_{\perp}^2)}{\alpha_s^2(p_{\perp}^2)} \frac{p_{\perp 0}^4}{(p_{\perp 0}^2 + p_{\perp}^2)^2} \quad (41.16)$$

would smoothly regulate the divergences, now with  $p_{\perp 0}$  as the free parameter. Regardless of whether it is imposed as a smooth (PYTHIA and SHERPA) or steep (HERWIG++) function, this is effectively the main “tuning” parameter in such models.

Note that the numerical value obtained for the cross section depends upon the PDF set used, and therefore the optimal value to use for the cutoff will also depend on this choice. Note also that the cutoff does not have to be energy-independent. Higher energies imply that parton densities can be probed at smaller  $x$  values, where the number of partons rapidly increases. Partons then become closer packed and the color screening distance  $d$  decreases. The uncertainty on the energy and/or  $x$  scaling of the cutoff is a major concern when extrapolating between different collider energies [87].

We now turn to the origin of the observational fact that hard jets appear to sit on top of a higher “pedestal” of underlying

activity than events with no hard jets. This is interpreted as a consequence of impact-parameter-dependence: in peripheral collisions, only a small fraction of events contain any high- $p_{\perp}$  activity, whereas central collisions are more likely to contain at least one hard scattering; a high- $p_{\perp}$  triggered sample will therefore be biased towards small impact parameters,  $b$ . The ability of a model to describe the shape of the pedestal (e.g. to describe both MB and UE distributions simultaneously) therefore depends upon its modeling of the  $b$ -dependence, and correspondingly the impact-parameter shape constitutes another main tuning parameter.

For each impact parameter  $b$ , the number of interactions  $\bar{n}(b)$  can still be assumed to be distributed according to Eq. (41.15), again modulo momentum conservation, but now with the mean value of the Poisson distribution depending on impact parameter,  $\langle \bar{n}(b) \rangle$ . This causes the final  $n$ -distribution (integrated over  $b$ ) to be wider than a Poissonian.

Finally, there are two perturbative modeling aspects which go beyond the introduction of MPI themselves: 1) parton showers off the MPI, and 2) perturbative parton-rescattering effects. Without showers, MPI models would generate very sharp peaks for back-to-back MPI jets, caused by unshowered partons passed directly to the hadronization model. However, with the exception of the oldest PYTHIA6 model, all GPMC models do include such showers [15], and hence should exhibit more realistic (i.e., broader and more decorrelated) MPI jets. On the initial-state side, the main questions are whether and how correlated multi-parton densities are taken into account and, as discussed previously, how the showers are regulated at low  $p_{\perp}$  and/or low  $x$ . Although none of the MC models currently impose a rigorous correlated multi-parton evolution, all of them include some elementary aspects. The most significant for parton-level results is arguably momentum conservation, which is enforced explicitly in all the models. The so-called “interleaved” models [24] attempt to go a step further, generating an explicitly correlated multi-parton evolution in which flavor sum rules are imposed to conserve, e.g. the total numbers of valence and sea quarks [60].

Perturbative rescattering in the final state can occur if partons are allowed to undergo several distinct interactions, with showering activity possibly taking place in-between. This has so far not been studied extensively, but a first exploratory model is available [88]. In the initial state, parton rescattering/recombination effects have so far not been included in any of the GPMC models.

#### 41.3.3. Bose-Einstein and Color-Reconnection Effects :

In the context of  $e^+e^-$  collisions, Bose-Einstein (BE) correlations have mostly been discussed as a source of uncertainty on high-precision  $W$  mass determinations at LEP [89]. In hadron-hadron (and nucleus-nucleus) collisions, however, BE correlations are used extensively to study the space-time structure of hadronizing matter (“femtoscopy”).

In MC models of hadronization, each string break or particle/cluster decay is normally factorized from all other ones. This reduces the number of variables that must be considered simultaneously, but also makes the introduction of correlations among particles from different breaks/decays intrinsically difficult to address. In the context of GPMCs, a few semi-classical models are available within the PYTHIA 6 and 8 generators [90], in which the BE effect is mimicked by an attractive interaction between pairs of identical particles in the final state, with no higher correlations included. This “force” acts after the decays of very short-lived particles, like  $\rho$ , but before decays of longer-lived ones, like  $\pi^0$ . The main differences between the variants of this model is the assumed shape of the correlation function and how overall momentum conservation is handled.

As discussed in Sec. 41.2, leading-color (“planar”) color flows are used to set up the hadronizing systems (clusters or strings) at the hadronization stage. If the systems do not overlap significantly in space and time, subleading-color ambiguities and/or non-perturbative reconnections are expected to be small. However, if the density of displaced color charges is sufficiently high that several systems can overlap significantly, full-color and/or reconnection effects should become progressively larger.

In the specific context of MPI, a crucial question is how color is neutralized *between* different MPI systems, including the remnants.

The large rapidity differences involved imply large invariant masses (though normally low  $p_{\perp}$ ), and hence large amounts of (soft) particle production. Indeed, in the context of soft-inclusive physics, it is these “inter-system” strings/clusters that furnish the dominant particle-production mechanism, and hence their modeling is an essential part of the soft-physics description, affecting topics such as MB/UE multiplicity and  $p_{\perp}$  distributions, rapidity gaps, and precision mass measurements. More comprehensive reviews of color-reconnection effects can be found in Refs. 15,91.

#### 41.4. Parameters and Tuning

The accuracy that can be achieved by a GPMC model depends on the sophistication of theory models it incorporates, on the inclusiveness of the observable(s) under study, and on the available constraints on its free parameters. Using existing data (or more precise theory calculations) to constrain the model parameters is referred to as generator tuning.

Typically, the most inclusive event properties are determined by only a few, very important parameters, such as the value of  $\alpha_S$ , for perturbative corrections, and the shape of the fragmentation functions, for non-perturbative ones. More parameters may then be introduced to describe successively more exclusive aspects, but these should have progressively less impact on the overall modelling. One may therefore take a factorized approach, first constraining the perturbative parameters and thereafter the non-perturbative ones, in order of decreasing significance to the overall modeling. Furthermore, by identifying which measurements are most sensitive to each parameter, this ordering can be reflected in the way that data is selected and applied to constrain the models. Thus, the most inclusive measurements should be used first, to constrain the most inclusive parameters, and so on for progressively more exclusive aspects.

At LO $\times$ LL, perturbation theory is doing well if it agrees with an IR safe measurement within 10%. It would therefore not make much sense to tune a GPMC beyond roughly 5% (it might even be dangerous, due to overfitting). The advent of NLO Monte Carlos may reduce this number slightly, but only for quantities for which one expects NLO precision. For LO Monte Carlos, distributions should be normalized to unity, since the NLO normalization is not tunable. For quantities governed by non-perturbative physics, uncertainties are larger. For some quantities, e.g. ones for which the underlying modeling is known to be poor, an order-of-magnitude agreement or worse may have to be accepted.

In the context of LO $\times$ LL GPMC tuning, subleading aspects of coupling-constant and PDF choices are relevant. In particular, one should be aware that the choice of QCD  $\Lambda$  parameter  $\Lambda_{\overline{\text{MS}}} = 1.569\Lambda_{\overline{\text{MS}}}$  (for 5 active flavors) improves the predictions of coherent shower algorithms at the NLL level [92], and hence this scheme is typically considered the baseline for shower tuning. The question of LO vs. NLO PDFs is more involved [15], but it should be emphasized that the low- $x$  gluon in particular is important for determining the level of the underlying event in MPI models (Sec. 41.3.2), and hence the MB/UE tuning (and energy scaling [87]) is linked to the choice of PDF in such models. Further issues and an example of a specific recipe that could be followed in a realistic set-up can be found in Ref. 93. A useful online resource can be found at the mcplots.cern.ch web site [94], based on the RIVET tool [95].

Recent years have seen the emergence of automated tools to reduce the amount of both computer and manpower required for tuning [96]. Automating the human expert input is more difficult. In the tools currently on the market, this is addressed by a combination of input solicited from the GPMC authors (e.g., which parameters and ranges to consider, which observables constitute a complete set, etc) and a set of weights determining the relative priority given to each bin in each distribution. Studies of sensitivities and correlations also play an important role. Overall, the quality of the resulting tunes is by now competitive. The field is still burgeoning, with future sophistications to be expected.

#### References

1. G. Corcella *et al.*, JHEP **0101**, 010 (2001), [hep-ph/0011363](#).
2. M.Bähr *et al.*, Eur. Phys. J. **C58**, 639 (2008), [arXiv:0803.0883](#).

3. T. Sjöstrand, S. Mrenna, and P. Z. Skands, JHEP **05**, 026 (2006), [hep-ph/0603175](#).
4. T. Sjöstrand *et al.*, Comp. Phys. Comm. **191**, 159 (2015), [arXiv:1410.3012](#).
5. T. Gleisberg *et al.*, JHEP **0402**, 056 (2004), [hep-ph/0311263](#).
6. T. Kinoshita, J. Math. Phys. **3**, 650 (1962).
7. T. Lee and M. Nauenberg, Phys. Rev. **133**, 1549 (1964).
8. G.C. Fox and S. Wolfram, Nucl. Phys. **B168**, 285 (1980).
9. G. Altarelli and G. Parisi, Nucl. Phys. **B126**, 298 (1977).
10. A. Einstein, B. Podolsky, and N. Rosen, Phys. Rev. **47**, 777 (1935).
11. B.R. Webber, Phys. Lett. **B193**, 91 (1987).
12. J.C. Collins, Nucl. Phys. **B304**, 794 (1988).
13. I.G. Knowles, Comp. Phys. Comm. **58**, 271 (1990).
14. T. Sjöstrand, Phys. Lett. **B157**, 321 (1985).
15. A. Buckley *et al.*, Phys. Reports **504**, 145 (2011), [arXiv:1101.2599](#).
16. G. Marchesini and B.R. Webber, Nucl. Phys. **B310**, 461 (1988).
17. S. Gieseke, P. Stephens, and B. Webber, JHEP **0312**, 045 (2003), [hep-ph/0310083](#).
18. M. Bengtsson and T. Sjöstrand, Nucl. Phys. **B289**, 810 (1987).
19. G. Gustafson and U. Pettersson, Nucl. Phys. **B306**, 746 (1988).
20. L. Lönnblad, Comp. Phys. Comm. **71**, 15 (1992).
21. W.T. Giele, D.A. Kosower, and P.Z. Skands, Phys. Rev. **D78**, 014026 (2008), [arXiv:0707.3652](#).
22. S. Schumann and F. Krauss, JHEP **0803**, 038 (2008), [arXiv:0709.1027](#).
23. Z. Nagy and D.E. Soper, JHEP **0510**, 024 (2005), [hep-ph/0503053](#).
24. T. Sjöstrand and P.Z. Skands, Eur. Phys. J. **C39**, 129 (2005), [hep-ph/0408302](#).
25. E. Norrbin and T. Sjöstrand, Nucl. Phys. **B603**, 297 (2001), [hep-ph/0010012](#).
26. S. Catani *et al.*, Nucl. Phys. **B627**, 189 (2002), [hep-ph/0201036](#).
27. J. Cembranos *et al.*, (2013), [arXiv:1305.2124](#).
28. N. Davidson, T. Przedzinski, and Z. Was, (2010), [arXiv:1011.0937](#).
29. K. Hamilton and P. Richardson, JHEP **0607**, 010 (2006), [hep-ph/0603034](#).
30. M. Schönherr and F. Krauss, JHEP **0812**, 018 (2008), [arXiv:0810.5071](#).
31. A. Semenov, Comp. Phys. Comm. **180**, 431 (2009), [arXiv:0805.0555](#).
32. N.D. Christensen and C. Duhr, Comp. Phys. Comm. **180**, 1614 (2009), [arXiv:0806.4194](#).
33. M. Fairbairn *et al.*, Phys. Reports **438**, 1 (2007), [hep-ph/0611040](#).
34. J. Alwall *et al.*, Comp. Phys. Comm. **176**, 300 (2007), [hep-ph/0609017](#).
35. P.Z. Skands *et al.*, JHEP **0407**, 036 (2004), [hep-ph/0311123](#).
36. J. Alwall *et al.*, (2007), [arXiv:0712.3311](#).
37. P. Richardson, JHEP **0111**, 029 (2001), [hep-ph/0110108](#).
38. M. Bengtsson and T. Sjöstrand, Phys. Lett. **B185**, 435 (1987).
39. W. T. Giele, D. A. Kosower and P. Z. Skands, Phys. Rev. **D84**, 054003 (2011), [arXiv:1102.2126](#).
40. S. Catani *et al.*, JHEP **11**, 063 (2001), [hep-ph/0109231](#).
41. J. Alwall *et al.*, Eur. Phys. J. **C53**, 473 (2008), [arXiv:0706.2569](#).
42. P. Nason, JHEP **11**, 040 (2004), [hep-ph/0409146](#).
43. B. Cooper *et al.*, Eur. Phys. J. **C72**, 2078 (2012), [arXiv:1109.5295](#).
44. S. Frixione and B.R. Webber, JHEP **06**, 029 (2002), [hep-ph/0204244](#).
45. S. Alioli *et al.*, JHEP **1006**, 043 (2010), [arXiv:1002.2581](#).
46. S. Alioli, K. Hamilton, and E. Re, JHEP **09**, 104 (2011), [arXiv:1108.0909](#).
47. R. Frederix and S. Frixione, JHEP **12**, 061 (2012), [arXiv:1209.6215](#).
48. L. Lönnblad and S. Prestel, JHEP **03**, 166 (2013), [arXiv:1211.7278](#).
49. K. Hamilton, P. Nason and G. Zanderighi, JHEP **10**, 155 (2012), [arXiv:1206.3572](#).

50. S. Höche *et al.*, JHEP **04**, 027 (2013), [arXiv:1207.5030](#).
51. K. Hamilton, P. Nason and G. Zanderighi, JHEP **05**, 140 (2015), [arXiv:1501.04637](#).
52. M.L. Mangano *et al.*, JHEP **0307**, 001 (2003), [hep-ph/0206293](#).
53. J. Alwall *et al.*, JHEP **1106**, 128 (2011), [arXiv:1106.0522](#).
54. E. Boos *et al.*, (2007), [hep-ph/0109068](#).
55. J. Alwall *et al.*, JHEP **07**, 079 (2014), [arXiv:1405.0301](#).
56. G. Gustafson, Phys. Lett. **B175**, 453 (1986).
57. X. Artru and G. Mennessier, Nucl. Phys. **B70**, 93 (1974).
58. B. Andersson *et al.*, Phys. Reports **97**, 31 (1983).
59. B. Andersson, Camb. Monogr. Part. Phys. Nucl. Phys. Cosmol. **7** (1997).
60. T. Sjöstrand and P.Z. Skands, JHEP **0403**, 053 (2004), [hep-ph/0402078](#).
61. M. Bowler, Z. Phys. **C11**, 169 (1981).
62. C. Peterson *et al.*, Phys. Rev. **D27**, 105 (1983).
63. D. Amati and G. Veneziano, Phys. Lett. **B83**, 87 (1979).
64. A. Bassetto, M. Ciafaloni, and G. Marchesini, Phys. Lett. **B83**, 207 (1979).
65. R.D. Field and S. Wolfram, Nucl. Phys. **B213**, 65 (1983).
66. B.R. Webber, Nucl. Phys. **B238**, 492 (1984).
67. J.-C. Winter, F. Krauss, and G. Soff, Eur. Phys. J. **C36**, 381 (2004), [hep-ph/0311085](#).
68. D. Lange, Nucl. Instrum. Methods **A462**, 152 (2001).
69. S. Jadach *et al.*, Comp. Phys. Comm. **76**, 361 (1993).
70. D. Grellscheid and P. Richardson, (2007), [arXiv:0710.1951](#).
71. T. Gleisberg *et al.*, JHEP **0902**, 007 (2009), [arXiv:0811.4622](#).
72. V. Khoze *et al.*, Eur. Phys. J. **C69**, 85 (2010), [arXiv:1005.4839](#).
73. T. Sjöstrand and M. van Zijl, Phys. Rev. **D36**, 2019 (1987).
74. J.M. Butterworth, J.R. Forshaw, and M.H. Seymour, Z. Phys. **C72**, 637 (1996), [hep-ph/9601371](#).
75. M. Bähr *et al.*, (2009), [arXiv:0905.4671](#).
76. R. Corke and T. Sjöstrand, JHEP **1105**, 009 (2011), [1101.5953](#).
77. G.A. Schuler and T. Sjöstrand, Phys. Rev. **D49**, 2257 (1994).
78. G. Ingelman and P. Schlein, Phys. Lett. **B152**, 256 (1985).
79. S. Navin, (2010), [arXiv:1005.3894](#).
80. P. Aurenche *et al.*, Comp. Phys. Comm. **83**, 107 (1994), [hep-ph/9402351](#).
81. F.W. Bopp, R. Engel, and J. Ranft, (1998), [hep-ph/9803437](#).
82. S. Roesler, R. Engel, and J. Ranft, p. 1033 (2000), [hep-ph/0012252](#).
83. B.E. Cox and J.R. Forshaw, Comp. Phys. Comm. **144**, 104 (2002), [hep-ph/0010303](#).
84. M. Ryskin, A. Martin, and V. Khoze, Eur. Phys. J. **C71**, 1617 (2011), [arXiv:1102.2844](#).
85. P. Skands, S. Carrazza and J. Rojo, Eur. Phys. J. **C74**, 3024 (2014), [arXiv:1404.5630](#).
86. M. Bähr, J.M. Butterworth, and M.H. Seymour, JHEP **01**, 065 (2009), [arXiv:0806.2949](#).
87. H. Schulz and P.Z. Skands, Eur. Phys. J. **C71**, 1644 (2011), [arXiv:1103.3649](#).
88. R. Corke and T. Sjöstrand, JHEP **01**, 035 (2009), [arXiv:0911.1901](#).
89. LEP Electroweak Working Group, (2005), [hep-ex/0511027](#).
90. L. Lönnblad and T. Sjöstrand, Eur. Phys. J. **C2**, 165 (1998), [hep-ph/9711460](#).
91. J. R. Christiansen and P. Z. Skands, JHEP **08**, 003 (2015), [arXiv:1505.01681](#).
92. S. Catani, B. R. Webber, and G. Marchesini, Nucl. Phys. **B349**, 635 (1991).
93. P.Z. Skands, (2011), [arXiv:1104.2863](#).
94. A. Karneyeu *et al.*, Eur. Phys. J. **C74**, 2714 (2013), [arXiv:1306.3436](#).
95. A. Buckley *et al.*, Comp. Phys. Comm. **184**, 2803 (2010), [arXiv:1003.0694](#).
96. A. Buckley *et al.*, Eur. Phys. J. **C65**, 331 (2010), [arXiv:0907.2973](#).

## 42. MONTE CARLO NEUTRINO EVENT GENERATORS

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Monte Carlo neutrino generators are programs or libraries which simulate neutrino interactions with electrons, nucleons and nuclei. In this capacity their usual task is to take an input neutrino and nucleus and produce a set of 4-vectors for particles emerging from the interaction, which are then input to full detector simulations. Since these generators have to simulate not only the initial interaction of neutrinos with target particles, but re-interactions of the generated particles in the nucleus, they contain a wide range of elementary particle and nuclear physics. Viewed more broadly, they are the access point for neutrino experimentalists to the theory inputs needed for analysis. Examples include cross section libraries for event rate calculations and parameter uncertainties and reweighting tools for systematic error evaluation.

Neutrino experiments typically operate in neutrino beams that are neither completely pure nor mono-energetic. Generators are a crucial component in the convolution of beam flux, neutrino interaction physics, and detector response that is necessary to make predictions about observable quantities. Similarly they are used to relate reconstructed quantities back to true quantities. In these various capacities they are used from the detector design stage through the extraction of physics measurements from reconstructed observables. Monte Carlo neutrino generators play unique and important roles in the experimental study of neutrino interactions and oscillations.

There are several neutrino event generators available, such as ANIS [1], GENIE [2], GiBUU [3], NEGN [4], NEUT [5], NUANCE [6], the FLUKA routines NUNDIS/NUNRES [7], and NuWRO [8]. Historically, experiments would develop their own generators. This was often because they were focused on a particular measurement, energy range, or target, and wanted to ensure that the best physics was included for it. These ‘home-grown’ generators were often tuned primarily or exclusively to the neutrino data most similar to the data that the experiment would be collecting. A major advance in the field was the introduction of conference series devoted to the topic of neutrino interaction physics, NuINT and NuFACT in particular. Event generator comparisons have been a regular staple of the NuINT conference series from its inception, and a great deal of information on this topic can be found in the Proceedings of these meetings. These meetings have facilitated experiment-theory discussions leading to the first generator developed by a theory group (NuWRO) [8], the extension of established nuclear interaction codes (FLUKA and GiBUU) to include neutrino-nuclear processes [3,7], and inclusion of theorists in existing generator development teams.

These activities have led to more careful scrutiny of the crucial nuclear theory inputs to these generators, which is evaluated in particular through comparisons to electron-scattering data. At this point in time all simulation codes face challenges in describing the full extent of the lepton scattering data, and the tension between incorporating the best available theory versus obtaining the best agreement with the data plays out in a variety of ways within the field. For the field to make progress, inclusion of state of the art theory needs to be coupled to global analyses that correctly incorporate correlations between measurements. Given the rapid pace of new data and the complexity of analyses, this is a significant challenge for the field in the coming years.

There are many neutrino experiments which use various sources of neutrinos, from reactors, accelerators, the atmosphere, and astrophysical sources, thereby covering a range of energies from MeV to TeV. Much of the emphasis has been on the few-GeV region in the generators, as this is the relevant energy range for long-baseline neutrino oscillation experiments. These generators use the impulse approximation for most of the primary neutrino interactions and simulate the interactions of secondary particles in the nucleus in semi-classical ways in order to simulate a variety of nuclei in a single model, and for practical considerations as these approaches are fast. However, there are several challenges facing these simulations coming mainly from the complexity of the nuclear physics, and avoiding double counting in combining perturbative and non-perturbative models for the neutrino-nucleon scattering processes. While generators share

many common ingredients, differences in implementation, parameter values, and approaches to avoid double counting can yield dramatically different predictions [9]. In the following sections, interaction models and their implementations including the interactions of generated particles in the nuclei are described.

In order to assure its reproducibility, neutrino event generators are tuned and validated against a wide variety of data, including data from photon, charged lepton, neutrino, and hadron probes. The results from these external data tuning exercises are important for experiments as they quantify the uncertainty on model parameters, needed by experiments in the evaluation of generator-related systematic errors. Electron scattering data plays an important role in determining the vector contribution to the form-factors and structure functions, as well as in evaluating specific aspects of the nuclear model. Hadron scattering data is used in validating the nuclear model, in particular the modeling of final state interactions. Tuning of neutrino-nucleon scattering and hadronization models relies heavily on the previous generation of high energy neutrino scattering and hydrogen and deuterium bubble chamber experiments, and more recent data from the K2K, MiniBooNE, NOMAD, SciBooNE, MINOS, T2K, ArgoNEUT, and MINERvA experiments either has been, or will be, used for this purpose.

### 42.1. Neutrino-Nucleon Scattering

Event generators typically begin with free-nucleon cross sections which are then embedded into a nuclear physics model. The most important processes are quasi-elastic (elastic for NC) scattering, resonance production, and non-resonant inelastic scattering, which make comparable contributions for few-GeV interactions. The neutrino cross sections in this energy range can be seen in Figures 50.1 through 50.4 of this *Review*.

**42.1.1. Quasi-Elastic Scattering :** The cross section for the neutrino nucleon charged current quasi-elastic scattering is described in terms of the leptonic and hadronic weak currents, where dominant contributions to the hadronic current come from the vector and axial-vector form factors. There also exists the pseudo-scalar term (the pseudo-scalar form factor) in the hadronic current but this term is rather small for electron and muon neutrinos and usually related to the axial form factor assuming partially conserved axial current (PCAC). The vector form factors are measured by the recent precise electron scattering experiments and known to have some deviation from the simple dipole form [10]. Therefore, most of the generators use parametrizations of this form factor taken directly from the data. For the axial form factor there is no such precise experiment, and most of the generators use a dipole form. Generally, the value of axial form factor at  $q^2 = 0$  is extracted from the polarized nucleon beta decay experiment. However, the selection of the axial vector mass parameter depends on each generator, with values typically around  $1.00 \text{ GeV}/c^2$ .

**42.1.2. Resonance Production :** Most generators use the calculation of Rein-Sehgal to simulate neutrino-induced single pion production [13]. To obtain the cross section for a particular channel, they calculate the amplitude for the production of each resonance multiplied by the probability for the decay of that resonance into that particular channel. Implementation differences include the number of resonances included, whether the amplitudes are added coherently or incoherently, the invariant mass range over which the model is used, how non-resonant backgrounds are included, inclusion of lepton mass terms, and the model parameter values (in particular the axial mass). In this model it is also possible to calculate the cross-sections of single photon, kaon and  $\eta$  productions by changing the decay probability of the resonances, which are included in some of the programs. However, it is known that discrepancies exist between the recent pion electro/photoproduction data and the results from the simulation data with the same framework, i.e. vector part of this model. There are several attempts to overcome this issue [12] and some of the generators started using more appropriate form factors. The GiBUU and NuWRO generators do not use the Rein-Sehgal model, and instead rely directly on electro-production data for the vector contribution and fit bubble chamber data to determine the remaining parameters for the axial contribution [14], [15,16].

**42.1.3. Deep and Shallow Inelastic Scattering :** For this process the fundamental target shifts from the nucleon to its quark constituents. Therefore, the generators use the standard expression for the constructions for the nucleon structure functions  $F_2$  and  $xF_3$  from parton distributions for high  $Q^2$  (the DIS regime) to calculate direction and momentum of lepton. The first challenge is in extending this picture to the lower values of  $Q^2$  and  $W$  that dominate the available phase space for few-GeV interactions (the so-called ‘shallow inelastic scattering’, or SIS regime). The corrections proposed in [17] are widely used, while others [7] implement their own modifications to the parton distributions at low  $Q^2$ . Both DIS and SIS generates hadrons but their production depends on each generator’s implementation of a hadronization model as described in the next section. There are various difficulties not only in the actual hadronization but the relation with the single meson production. It is necessary to avoid double counting between the resonance and SIS/DIS models, and all generators are different in this regard. The scheme chosen can have a significant impact on the results of simulations at a few-GeV neutrino energies.

## 42.2. Hadronization Models

For hadrons produced via baryonic resonances, the underlying model amplitudes and resonance branching fractions can be used to fully characterize the hadronic system. For non-resonant production, a hadronization model is required. Most generators use PYTHIA [18] for this purpose, although some with modified parameters. In addition some implement their own models to handle invariant masses that are too low for PYTHIA, typically somewhere around 2.0 GeV/ $c^2$ . Such models rely heavily on measurements of neutrino hadro-production in high-resolution devices, such as bubble chambers and the CHORUS [19] and NOMAD experiments [20], to construct empirical parametrizations that reproduce the key features of the data [21,22]. The basic ingredients are the empirical observations that average charged particle multiplicities increase logarithmically with the invariant mass of the hadronic system, and that the distribution of charged particle multiplicities about this average are described by a single function (an observation known as KNO scaling). Neutral particles are assumed to be produced with an average multiplicity that is 50% of the charged particle multiplicity. Simple parametrizations to more accurately reproduce differences observed in the forward/backward hemispheres of hadronic systems are included in GENIE, NEUT, and NuWRO.

## 42.3. Nuclear Physics

The nuclear physics relevant to neutrino-nucleus scattering at few-GeV energies is complicated, involving Fermi motion, nuclear binding, Pauli blocking, in-medium modifications of form factors and hadronization, intranuclear rescattering of hadrons, and many-body scattering mechanisms including long- and short-range nucleon-nucleon correlations.

### 42.3.1. Scattering Mechanisms :

Most of the models used for neutrino-nuclear scattering kinematics were developed in the context of few-GeV inclusive electron scattering, by experiments going back nearly 50 years. A topic of considerable discussion within this community has been to what extent the impulse approximation, whereby the nucleus is envisioned as collection of bound, moving, single nucleons, is appropriate. The question arose initially in the context of measurements of the quasi-elastic axial mass, with a number of recent experiments using nuclear targets measuring values that were significantly higher than those obtained by an earlier generation of bubble chamber experiments using hydrogen or deuterium [23]. These led to a revisitation of the role played by scattering from multi-particle/hole states in the nucleus. The contribution of these scattering processes is an extremely active area of theoretical research at present, with significant implications for generators and analyses [24]. The GiBUU, NuWRO, GENIE, and NEUT generators have all implemented, or are in the process of implementing, first models for these processes [25].

In order to obtain the cross-section off nucleons in the nucleus, it is necessary to take into account the in-medium effects. The basic

models employed in event generators rely on impulse approximation schemes, the most simple of which is the Relativistic Fermi Gas Model. The most common implementations are the Smith-Moniz [26] and Bodek-Ritchie [27] models. Within the electron scattering community, the analogous calculations have for decades relied on spectral functions, which incorporate information about nucleon momenta and binding energies in the impulse approximation scheme. The NuWRO and GiBUU generators currently use spectral functions, they are incorporated into NEUT as an option, and several of the other generators are incorporating spectral function models at this time. It is known from photo and electro-nuclear scattering that the Delta width is affected by Pauli blocking and collisional broadening. These effects are included in some, but not all, generators.

When scattering from a nucleus, coherent scattering of various kinds is possible. Most simulations incorporate, at least, neutral and charged coherent single pion production. While the interaction rate for these interactions is typically around a percent of the total yield, the unique kinematic features of these events can make them potential backgrounds for oscillation searches. Implemented in Monte Carlo are PCAC-based methods, while microscopic models are currently being incorporated into several generators as well. Reference [9] clearly demonstrates a point mentioned earlier, where generators implementing the same model [28] are seen to produce very different predictions.

### 42.3.2. Hadron Production in Nuclei :

Neutrino pion production is one of the dominant interactions in a few-GeV region and the interaction cross sections of pions in nucleus from those interactions are quite large. Therefore, the interactions of pions in nucleus changes the kinematics of the pions and can have large effects on the results of simulations at these energies. Most generators implement this physics through an intranuclear cascade simulation. In generators which utilize cascade models, a hadron, which has been formed in the nucleus, is moved step by step until it interacts with the other nucleon or escapes from the nucleus. The probabilities of each interaction in nucleus are usually given as the mean free paths and used to determine whether the hadron is interacted or not. If the hadron is found to be interacted, appropriate interactions are selected and simulated. Usually, absorption, elastic, and inelastic scatterings including particle productions are simulated as secondary interactions. The determination method of the kinematics for the final state particles heavily depends on the generators but most of them use experimentally validated models to simulate hadron interactions in nucleus. No two interanuclear cascade simulations implemented in neutrino event generators are the same. In all cases hadrons propagate from an interaction vertex chosen based on the density distribution of the target nucleus. In determining the generated position of the hadrons in nucleus, the concept of the formation length is sometimes employed. Based on this idea, the hadronization process is not instantaneous and it takes some time before generating the hadrons [29]. The basis for formation times are measurements at relatively high energy and  $Q^2$ , and most generators that employ the concept do not apply them to resonance interactions, the exception is [29]. The intranuclear rescattering simulations are typically validated against hadron scattering data. In some simulations (e.g. NEUT) the pion-less Delta decay is also considered and 20% of the events do not have a pion and only the lepton and the nucleon are generated.

The exception is GiBUU, a semiclassical transport model in coupled channels that describes the space-time evolution of a manybody system in the presence of potentials and a collision term [3]. This approach assures consistency between nuclear effects in the initial state, such as Fermi motion, Pauli blocking, hadron self-energies, and modified cross sections, and the final state, such as particle reinteractions, since the two are derived from the same model. This model has been previously used to describe a wide variety of nuclear interaction data. Similarly, the hadronic simulation of the NUNDIS/NUNRES programs are handled by the well-established FLUKA hadronic simulation package [7].

## References:

1. A. Gazizov and M. P. Kowalski, *Comp. Phys. Comm.* **172**, 203 (2005) [[astro-ph/0406439](#)].
2. C. Andreopoulos *et al.*, *Nucl. Instrum. Methods* **A614**, 87 (2010) [[arXiv:0905.2517](#)].
3. O. Buss *et al.*, *Phys. Reports* **512**, 1 (2012) [[arXiv:1106.1344](#)].
4. D. Autiero, *Nucl. Phys. (Proc. Supp.)* **139**, 253 (2005).
5. Y. Hayato, *Nucl. Phys. (Proc. Supp.)* **112**, 171 (2002).
6. D. Casper, *Nucl. Phys. (Proc. Supp.)* **112**, 161 (2002) [[hep-ph/0208030](#)].
7. G. Battiston *et al.*, *Acta Phys. Polon.* **B40**, 2491 (2009).
8. C. Juszczak, J.A. Nowak, and J.T. Sobczyk, *Nucl. Phys. (Proc. Supp.)* **159**, 211 (2006) [[hep-ph/0512365](#)].
9. S. Boyd *et al.*, *AIP Conf. Proc.* **1189**, 60 (2009).
10. A. Bodek *et al.*, *Eur. Phys. J.* **C53**, 349 (2008) [[arXiv:0708.1946](#)].
11. R.P. Feynman, M. Kislinger, and F. Ravndal, *Phys. Rev.* **D3**, 2706 (1971).
12. K.M. Graczyk and J.T. Sobczyk, *Phys. Rev.* **D77**, 053001 (2008); Erratum-ibid. **D79**, 079903 (2009) [[arXiv:0707.3561](#)].
13. D. Rein, and L.M. Sehgal, *Ann. Phys.* **133**, 79 (1981).
14. O. Lalakulich and E.A. Paschos, *Phys. Rev.* **D71**, 074003 (2005) [[hep-ph/0501109](#)].
15. J. Nowak, *Phys. Scripta* **T127**, 70 (2006).
16. L. Alvarez-Ruso, S.K. Singh, and M.J. Vicente-Vacas, *Phys. Rev.* **C57**, 2693 (1998) [[nucl-th/9712058](#)].
17. A. Bodek and U.K. Yang, *J. Phys.* **G29**, 1899 (2003) [[hep-ex/0210024](#)].
18. T. Sjostrand, S. Mrenna, and P. Skands, *JHEP* **0605**, 026 (2006) [[hep-ph/0603175](#)].
19. A. Kayis-Topaksu *et al.*, *Eur. Phys. J.* **C51**, 775 (2007) [[arXiv:0707.1586](#)].
20. J. Altegoer *et al.*, *Phys. Lett.* **B445**, 439 (1999).
21. T. Yang *et al.*, *Eur. Phys. J.* **C63**, 1 (2009) [[hep-ph/0904.4043](#)].
22. J. Nowak and J. Sobczyk, *Acta Phys. Polon.* **B37**, 2371 (2006) [[hep-ph/0608108](#)].
23. H. Gallagher, G. Garvey, G. Zeller, *Ann. Rev. Nucl. and Part. Sci.* **61**, 355 (2011).
24. O. Lalakulich, U. Mosel, and K. Gallmeister, *Phys. Rev.* **C86**, 054606 (2012) [[arXiv:1208.3678](#)].
25. T. Katori, *AIP Conf. Proc.* **1663**, 030001 (2015) [[arXiv:1304.6014](#)].
26. R. Smith and E. Moniz, *Nucl. Phys.* **B43**, 605 (1972).
27. A. Bodek and J. Ritchie, *Phys. Rev.* **D24**, 1400 (1981).
28. D. Rein and L. Sehgal, *Nucl. Phys.* **B223**, 29 (1983).
29. T. Golan, C. Juszczak, and J. Sobczyk, *Phys. Rev.* **C86**, 015505 (2012) [[arXiv:1202.4197](#)].



## 43. MONTE CARLO PARTICLE NUMBERING SCHEME

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The Monte Carlo particle numbering scheme presented here is intended to facilitate interfacing between event generators, detector simulators, and analysis packages used in particle physics. The numbering scheme was introduced in 1988 [1] and a revised version [2,3] was adopted in 1998 in order to allow systematic inclusion of quark model states which are as yet undiscovered and hypothetical particles such as SUSY particles. The numbering scheme is used in several event generators, *e.g.* HERWIG, PYTHIA, and SHERPA, and interfaces, *e.g.* /HEPEVT/ and HepMC.

The general form is a 7-digit number:

$$\pm n \ n_r \ n_L \ n_{q1} \ n_{q2} \ n_{q3} \ n_J.$$

This encodes information about the particle's spin, flavor content, and internal quantum numbers. The details are as follows:

1. Particles are given positive numbers, antiparticles negative numbers. The PDG convention for mesons is used, so that  $K^+$  and  $B^+$  are particles.
2. Quarks and leptons are numbered consecutively starting from 1 and 11 respectively; to do this they are first ordered by family and within families by weak isospin.
3. In composite quark systems (diquarks, mesons, and baryons)  $n_{q1-3}$  are quark numbers used to specify the quark content, while the rightmost digit  $n_J = 2J + 1$  gives the system's spin (except for the  $K_S^0$  and  $K_L^0$ ). The scheme does not cover particles of spin  $J > 4$ .
4. Diquarks have 4-digit numbers with  $n_{q1} \geq n_{q2}$  and  $n_{q3} = 0$ .
5. The numbering of mesons is guided by the nonrelativistic ( $L$ - $S$  decoupled) quark model, as listed in Tables 15.2 and 15.3.
  - a. The numbers specifying the meson's quark content conform to the convention  $n_{q1} = 0$  and  $n_{q2} \geq n_{q3}$ . The special case  $K_L^0$  is the sole exception to this rule.
  - b. The quark numbers of flavorless, light ( $u, d, s$ ) mesons are: 11 for the member of the isotriplet ( $\pi^0, \rho^0, \dots$ ), 22 for the lighter isosinglet ( $\eta, \omega, \dots$ ), and 33 for the heavier isosinglet ( $\eta', \phi, \dots$ ). Since isosinglet mesons are often large mixtures of  $u\bar{u} + d\bar{d}$  and  $s\bar{s}$  states, 22 and 33 are assigned by mass and do not necessarily specify the dominant quark composition.
  - c. The special numbers 310 and 130 are given to the  $K_S^0$  and  $K_L^0$  respectively.
  - d. The fifth digit  $n_L$  is reserved to distinguish mesons of the same total ( $J$ ) but different spin ( $S$ ) and orbital ( $L$ ) angular momentum quantum numbers. For  $J > 0$  the numbers are:  $(L, S) = (J-1, 1)$   $n_L = 0$ ,  $(J, 0)$   $n_L = 1$ ,  $(J, 1)$   $n_L = 2$  and  $(J+1, 1)$   $n_L = 3$ . For the exceptional case  $J = 0$  the numbers are  $(0, 0)$   $n_L = 0$  and  $(1, 1)$   $n_L = 1$  (*i.e.*  $n_L = L$ ). See Table 43.1.

**Table 43.1:** Meson numbering logic. Here  $qq$  stands for  $n_{q2} n_{q3}$ .

	$L = J - 1, S = 1$	$L = J, S = 0$	$L = J, S = 1$	$L = J + 1, S = 1$
$J$	code $J^{PC}$ $L$	code $J^{PC}$ $L$	code $J^{PC}$ $L$	code $J^{PC}$ $L$
0	— — —	00qq1 0 <sup>++</sup> 0	— — —	10qq1 0 <sup>++</sup> 1
1	00qq3 1 <sup>--</sup> 0	10qq3 1 <sup>+-</sup> 1	20qq3 1 <sup>++</sup> 1	30qq3 1 <sup>--</sup> 2
2	00qq5 2 <sup>++</sup> 1	10qq5 2 <sup>+-</sup> 2	20qq5 2 <sup>--</sup> 2	30qq5 2 <sup>++</sup> 3
3	00qq7 3 <sup>--</sup> 2	10qq7 3 <sup>+-</sup> 3	20qq7 3 <sup>++</sup> 3	30qq7 3 <sup>--</sup> 4
4	00qq9 4 <sup>++</sup> 3	10qq9 4 <sup>+-</sup> 4	20qq9 4 <sup>--</sup> 4	30qq9 4 <sup>++</sup> 5

- e. If a set of physical mesons correspond to a (non-negligible) mixture of basis states, differing in their internal quantum numbers, then the lightest physical state gets the smallest basis state number. For example the  $K_1(1270)$  is numbered 10313 ( $1^1 P_1 K_{1B}$ ) and the  $K_1(1400)$  is numbered 20313 ( $1^3 P_1 K_{1A}$ ).
- f. The sixth digit  $n_r$  is used to label mesons radially excited above the ground state.
- g. Numbers have been assigned for complete  $n_r = 0$   $S$ - and  $P$ -wave multiplets, even where states remain to be identified.
- h. In some instances assignments within the  $q\bar{q}$  meson model are only tentative; here best guess assignments are made.

i. Many states appearing in the Meson Listings are not yet assigned within the  $q\bar{q}$  model. Here  $n_{q2-3}$  and  $n_J$  are assigned according to the state's likely flavors and spin; all such unassigned light isoscalar states are given the flavor code 22. Within these groups  $n_L = 0, 1, 2, \dots$  is used to distinguish states of increasing mass. These states are flagged using  $n = 9$ . It is to be expected that these numbers will evolve as the nature of the states are elucidated. Codes are assigned to all mesons which are listed in the one-page table at the end of the Meson Summary Table as long as they have a preferred or established spin. Additional heavy meson states expected from heavy quark spectroscopy are also assigned codes.

6. The numbering of baryons is again guided by the nonrelativistic quark model, see Table 15.6. This numbering scheme is illustrated through a few examples in Table 43.2.
  - a. The numbers specifying a baryon's quark content are such that in general  $n_{q1} \geq n_{q2} \geq n_{q3}$ .
  - b. Two states exist for  $J = 1/2$  baryons containing 3 different types of quarks. In the lighter baryon ( $\Lambda, \Xi, \Omega, \dots$ ) the light quarks are in an antisymmetric ( $J = 0$ ) state while for the heavier baryon ( $\Sigma^0, \Xi', \Omega', \dots$ ) they are in a symmetric ( $J = 1$ ) state. In this situation  $n_{q2}$  and  $n_{q3}$  are reversed for the lighter state, so that the smaller number corresponds to the lighter baryon.
  - c. For excited baryons a scheme is adopted, where the  $n_r$  label is used to denote the excitation bands in the harmonic oscillator model, see Sec. 15.4. Using the notation employed there,  $n_r$  is given by the  $N$ -index of the  $D_N$  band identifier.
  - d. Further degeneracies of excited hadron multiplets with the same excitation number  $n_r$  and spin  $J$  are lifted by labelling such multiplets with the  $n_L$  index according to their mass, as given by its  $N$  or  $\Delta$ -equivalent.
  - e. In such excited multiplets extra singlets may occur, the  $\Lambda(1520)$  being a prominent example. In such cases the ordering is reversed such that the heaviest quark label is pushed to the last position:  $n_{q3} > n_{q1} > n_{q2}$ .
  - f. For pentaquark states  $n = 9$ ,  $n_r n_L n_{q1} n_{q2}$  gives the four quark numbers in order  $n_r \geq n_L \geq n_{q1} \geq n_{q2}$ ,  $n_{q3}$  gives the antiquark number, and  $n_J = 2J + 1$ , with the assumption that  $J = 1/2$  for the states currently reported.
7. The gluon, when considered as a gauge boson, has official number 21. In codes for glueballs, however, 9 is used to allow a notation in close analogy with that of hadrons.
8. The pomeron and odderon trajectories and a generic reggeon trajectory of states in QCD are assigned codes 990, 9990, and 110 respectively, where the final 0 indicates the indeterminate nature of the spin, and the other digits reflect the expected "valence" flavor content. We do not attempt a complete classification of all reggeon trajectories, since there is currently no need to distinguish a specific such trajectory from its lowest-lying member.
9. Two-digit numbers in the range 21–30 are provided for the Standard Model gauge bosons and Higgs.
10. Codes 81–100 are reserved for generator-specific pseudoparticles and concepts. Codes 901–920 are for additional non-standardized components of parton distribution functions.
11. The search for physics beyond the Standard Model is an active area, so these codes are also standardized as far as possible.
  - a. A standard fourth generation of fermions is included by analogy with the first three.
  - b. The graviton and the boson content of a two-Higgs-doublet scenario and of additional  $SU(2) \times U(1)$  groups are found in the range 31–40.
  - c. "One-of-a-kind" exotic particles are assigned numbers in the range 41–80.
  - d. Fundamental supersymmetric particles are identified by adding a nonzero  $n$  to the particle number. The superpartner of a boson or a left-handed fermion has  $n = 1$  while the superpartner of a right-handed fermion has  $n = 2$ . When mixing occurs, such as between the winos and charged Higgsinos to give charginos, or between left and right sfermions, the lighter physical state is given the smaller basis state number.

**Table 43.2:** Some examples of octet (top) and decuplet (bottom) members for the numbering scheme for excited baryons. Here  $qqq$  stands for  $n_{q_1}n_{q_2}n_{q_3}$ . See the text for the definition of the notation. The numbers in parenthesis correspond to the mass of the baryons. The states marked as (?) are not experimentally confirmed.

$J^P$	$(D, L_N^P)$	$n_r n_L n_{q_1} n_{q_2} n_{q_3} n_J$	$N$	$\Lambda_8$	$\Sigma$	$\Xi$	$\Lambda_1$
Octet			211,221	312	311,321,322	331,332	213
$1/2^+$	$(\mathbf{56}, \mathbf{0}_0^+)$	$00qqq2$	(939)	(1116)	(1193)	(1318)	—
$1/2^+$	$(\mathbf{56}, \mathbf{0}_2^+)$	$20qqq2$	(1440)	(1600)	(1660)	(1690)	—
$1/2^+$	$(\mathbf{70}, \mathbf{0}_2^+)$	$21qqq2$	(1710)	(1810)	(1880)	(?)	(?)
$1/2^-$	$(\mathbf{70}, \mathbf{1}_1^-)$	$10qqq2$	(1535)	(1670)	(1620)	(1750)	(1405)
$J^P$	$(D, L_N^P)$	$n_r n_L n_{q_1} n_{q_2} n_{q_3} n_J$	$\Delta$		$\Sigma$	$\Xi$	$\Omega$
Decuplet			111,211,221,222		311,321,322	331,332	333
$3/2^+$	$(\mathbf{56}, \mathbf{0}_0^+)$	$00qqq4$	(1232)		(1385)	(1530)	(1672)
$3/2^+$	$(\mathbf{56}, \mathbf{0}_2^+)$	$20qqq4$	(1600)		(1690)	(?)	(?)
$1/2^-$	$(\mathbf{70}, \mathbf{1}_1^-)$	$11qqq2$	(1620)		(1750)	(?)	(?)
$3/2^-$	$(\mathbf{70}, \mathbf{1}_1^-)$	$12qqq4$	(1700)		(?)	(?)	(?)

- e. Technicolor states have  $n = 3$ , with technifermions treated like ordinary fermions. States which are ordinary color singlets have  $n_r = 0$ . Color octets have  $n_r = 1$ . If a state has non-trivial quantum numbers under the topcolor groups  $SU(3)_1 \times SU(3)_2$ , the quantum numbers are specified by  $\text{tech}, ij$ , where  $i$  and  $j$  are 1 or 2.  $n_L$  is then  $2i + j$ . The coloron,  $V_8$ , is a heavy gluon color octet and thus is 3100021.
- f. Excited (composite) quarks and leptons are identified by setting  $n = 4$  and  $n_r = 0$ .
- g. Within several scenarios of new physics, it is possible to have colored particles sufficiently long-lived for color-singlet hadronic states to form around them. In the context of supersymmetric scenarios, these states are called  $R$ -hadrons, since they carry odd  $R$ -parity.  $R$ -hadron codes, defined here, should be viewed as templates for corresponding codes also in other scenarios, for any long-lived particle that is either an unflavored color octet or a flavored color triplet. The  $R$ -hadron code is obtained by combining the SUSY particle code with a code for the light degrees of freedom, with as many intermediate zeros removed from the former as required to make place for the latter at the end. (To exemplify, a sparticle  $n00000n_{\tilde{q}}$  combined with quarks  $q_1$  and  $q_2$  obtains code  $n00n_{\tilde{q}}n_{q_1}n_{q_2}n_J$ .) Specifically, the new-particle spin decouples in the limit of large masses, so that the final  $n_J$  digit is defined by the spin state of the light-quark system alone. An appropriate number of  $n_q$  digits is used to define the ordinary-quark content. As usual, 9 rather than 21 is used to denote a gluon/gluino in composite states. The sign of the hadron agrees with that of the constituent new particle (a color triplet) where there is a distinct new antiparticle, and else is defined as for normal hadrons. Particle names are  $R$  with the flavor content as lower index.
- h. A black hole in models with extra dimensions has code 5000040. Kaluza-Klein excitations in models with extra dimensions have  $n = 5$  or  $n = 6$ , to distinguish excitations of left- or right-handed fermions or, in case of mixing, the lighter or heavier state (cf. 11d). The nonzero  $n_r$  digit gives the radial excitation number, in scenarios where the level spacing allow these to be distinguished. Should the model also contain supersymmetry, excited SUSY states would be denoted by an  $n_r > 0$ , with  $n = 1$  or 2 as usual. Should some colored states be long-lived enough that hadrons would form around them, the coding strategy of 11g applies, with the initial two  $nn_r$  digits preserved in the combined code.
- i. Magnetic monopoles and dyons are assumed to have one unit of Dirac monopole charge and a variable integer number  $n_{q_1}n_{q_2}n_{q_3}$  units of electric charge. Codes  $411n_{q_1}n_{q_2}n_{q_3}0$  are then used when the magnetic and electrical charge sign agree and  $412n_{q_1}n_{q_2}n_{q_3}0$  when they disagree, with the overall sign

of the particle set by the magnetic charge. For now no spin information is provided.

- j. The nature of Dark Matter (DM) is not known, and therefore a definitive classification is too early. Candidates within specific scenarios are classified therein, such as 1000022 for the lightest neutralino. Generic fundamental states can be given temporary codes in the range 51 - 60, with 51, 52 and 53 reserved for spin 0, 1/2 and 1 ones. Generic mediators of s-channel DM pair creation or annihilation can be given codes 54 and 55 for spin 0 or 1 ones. Separate antiparticles, with negative codes, may or may not exist. More elaborate new scenarios should be constructed with  $n = 5$  and  $n_r = 9$ .
- k. Hidden Valley particles have  $n = 4$  and  $n_r = 9$ , and trailing numbers in agreement with their nearest-analog standard particles, as far as possible. Thus 4900021 is the gauge boson  $g_v$  of a confining gauge field, 490000 $n_{q_v}$  and 490001 $n_{\ell_v}$  fundamental constituents charged or not under this, 4900022 is the  $\gamma_v$  of a non-confining field, and 4900 $n_{q_v1}n_{q_v2}n_J$  a Hidden Valley meson.
12. Occasionally program authors add their own states. To avoid confusion, these should be flagged by setting  $nn_r = 99$ .
13. Concerning the non-99 numbers, it may be noted that only quarks, excited quarks, squarks, and diquarks have  $n_{q_3} = 0$ ; only diquarks, baryons (including pentaquarks), and the odderon have  $n_{q_1} \neq 0$ ; and only mesons, the reggeon, and the pomeron have  $n_{q_1} = 0$  and  $n_{q_2} \neq 0$ . Concerning mesons (not antimesons), if  $n_{q_1}$  is odd then it labels a quark and an antiquark if even.
14. Nuclear codes are given as 10-digit numbers  $\pm 10LZZZAAAI$ . For a (hyper)nucleus consisting of  $n_p$  protons,  $n_n$  neutrons and  $n_\Lambda$   $\Lambda$ 's,  $A = n_p + n_n + n_\Lambda$  gives the total baryon number,  $Z = n_p$  the total charge and  $L = n_\Lambda$  the total number of strange quarks.  $I$  gives the isomer level, with  $I = 0$  corresponding to the ground state and  $I > 0$  to excitations, see [4], where states denoted  $m, n, p, q$  translate to  $I = 1 - 4$ . As examples, the deuteron is 1000010020 and  $^{235}\text{U}$  is 1000922350. To avoid ambiguities, nuclear codes should not be applied to a single hadron, like  $p$ ,  $n$  or  $\Lambda^0$ , where quark-contents-based codes already exist.

This text and full lists of particle numbers can be found online [5].

#### References:

1. G.P. Yost *et al.*, Particle Data Group, Phys. Lett. **B204**, 1 (1988).
2. I.G. Knowles *et al.*, CERN 96-01, v. 2, p. 103.
3. C. Caso *et al.*, Particle Data Group, Eur. Phys. J. **C3**, 1 (1998).
4. G. Audi *et al.*, Nucl. Phys. **A729**, 3 (2003).
5. <http://pdg.lbl.gov/current/mc-particle-id/>.

**QUARKS**

$d$	1
$u$	2
$s$	3
$c$	4
$b$	5
$t$	6
$b'$	7
$t'$	8

**LEPTONS**

$e^-$	11
$\nu_e$	12
$\mu^-$	13
$\nu_\mu$	14
$\tau^-$	15
$\nu_\tau$	16
$\tau'^-$	17
$\nu_{\tau'}$	18

**GAUGE AND  
HIGGS BOSONS**

$g$	(9) 21
$\gamma$	22
$Z^0$	23
$W^+$	24
$h^0/H_1^0$	25
$Z'/Z_2^0$	32
$Z''/Z_3^0$	33
$W'/W_2^+$	34
$H^0/H_2^0$	35
$A^0/H_3^0$	36
$H^+$	37

**SPECIAL  
PARTICLES**

$G$ (graviton)	39
$R^0$	41
$LQ^c$	42
<b>DM(S = 0)</b>	<b>51*</b>
<b>DM(S = 1/2)</b>	<b>52*</b>
<b>DM(S = 1)</b>	<b>53*</b>
<i>reggeon</i>	110
<i>pomeron</i>	990
<i>odderon</i>	9990

for MC internal  
use 81–100 and 901–920

**DIQUARKS**

$(dd)_1$	1103
$(ud)_0$	2101
$(ud)_1$	2103
$(uu)_1$	2203
$(sd)_0$	3101
$(sd)_1$	3103
$(su)_0$	3201
$(su)_1$	3203
$(ss)_1$	3303
$(cd)_0$	4101
$(cd)_1$	4103
$(cu)_0$	4201
$(cu)_1$	4203
$(cs)_0$	4301
$(cs)_1$	4303
$(cc)_1$	4403
$(bd)_0$	5101
$(bd)_1$	5103
$(bu)_0$	5201
$(bu)_1$	5203
$(bs)_0$	5301
$(bs)_1$	5303
$(bc)_0$	5401
$(bc)_1$	5403
$(bb)_1$	5503

**SUSY****PARTICLES**

$\tilde{d}_L$	1000001
$\tilde{u}_L$	1000002
$\tilde{s}_L$	1000003
$\tilde{c}_L$	1000004
$\tilde{b}_1$	1000005 <sup>a</sup>
$\tilde{t}_1$	1000006 <sup>a</sup>
$\tilde{e}_L$	1000011
$\tilde{\nu}_{eL}$	1000012
$\tilde{\mu}_L$	1000013
$\tilde{\nu}_{\mu L}$	1000014
$\tilde{\tau}_1$	1000015 <sup>a</sup>
$\tilde{\nu}_{\tau L}$	1000016
$\tilde{d}_R$	2000001
$\tilde{u}_R$	2000002
$\tilde{s}_R$	2000003
$\tilde{c}_R$	2000004
$\tilde{b}_2$	2000005 <sup>a</sup>
$\tilde{t}_2$	2000006 <sup>a</sup>
$\tilde{e}_R$	2000011
$\tilde{\mu}_R$	2000013
$\tilde{\tau}_2$	2000015 <sup>a</sup>
$\tilde{g}$	1000021
$\tilde{\chi}_1^0$	1000022 <sup>b</sup>
$\tilde{\chi}_2^0$	1000023 <sup>b</sup>
$\tilde{\chi}_1^+$	1000024 <sup>b</sup>
$\tilde{\chi}_3^0$	1000025 <sup>b</sup>
$\tilde{\chi}_4^0$	1000035 <sup>b</sup>
$\tilde{\chi}_2^+$	1000037 <sup>b</sup>
$\tilde{G}$	1000039

**LIGHT  $I = 1$  MESONS**

$\pi^0$	111
$\pi^+$	211
$a_0(980)^0$	9000111
$a_0(980)^+$	9000211
$\pi(1300)^0$	100111
$\pi(1300)^+$	100211
$a_0(1450)^0$	10111
$a_0(1450)^+$	10211
$\pi(1800)^0$	9010111
$\pi(1800)^+$	9010211
$\rho(770)^0$	113
$\rho(770)^+$	213
$b_1(1235)^0$	10113
$b_1(1235)^+$	10213
$a_1(1260)^0$	20113
$a_1(1260)^+$	20213
$\pi_1(1400)^0$	9000113
$\pi_1(1400)^+$	9000213
$\rho(1450)^0$	100113
$\rho(1450)^+$	100213
$\pi_1(1600)^0$	9010113
$\pi_1(1600)^+$	9010213
$a_1(1640)^0$	9020113
$a_1(1640)^+$	9020213
$\rho(1700)^0$	30113
$\rho(1700)^+$	30213
$\rho(1900)^0$	9030113
$\rho(1900)^+$	9030213
$\rho(2150)^0$	9040113
$\rho(2150)^+$	9040213
$a_2(1320)^0$	115
$a_2(1320)^+$	215
$\pi_2(1670)^0$	10115
$\pi_2(1670)^+$	10215
$a_2(1700)^0$	9000115
$a_2(1700)^+$	9000215
$\pi_2(2100)^0$	9010115
$\pi_2(2100)^+$	9010215
$\rho_3(1690)^0$	117
$\rho_3(1690)^+$	217
$\rho_3(1990)^0$	9000117
$\rho_3(1990)^+$	9000217
$\rho_3(2250)^0$	9010117
$\rho_3(2250)^+$	9010217
$a_4(2040)^0$	119
$a_4(2040)^+$	219

**LIGHT  $I = 0$  MESONS**

$(u\bar{u}, d\bar{d}, \text{ and } s\bar{s} \text{ Admixtures})$	
$\eta$	221
$\eta'(958)$	331
$f_0(600)$	9000221
$f_0(980)$	9010221
$\eta(1295)$	100221
$f_0(1370)$	10221
$\eta(1405)$	9020221
$\eta(1475)$	100331
$f_0(1500)$	9030221
$f_0(1710)$	10331
$\eta(1760)$	9040221
$f_0(2020)$	9050221
$f_0(2100)$	9060221
$f_0(2200)$	9070221
$\eta(2225)$	9080221
$\omega(782)$	223
$\phi(1020)$	333
$h_1(1170)$	10223
$f_1(1285)$	20223
$h_1(1380)$	10333
$f_1(1420)$	20333
$\omega(1420)$	100223
$f_1(1510)$	9000223
$h_1(1595)$	9010223
$\omega(1650)$	30223
$\phi(1680)$	100333
$f_2(1270)$	225
$f_2(1430)$	9000225
$f_2'(1525)$	335
$f_2(1565)$	9010225
$f_2(1640)$	9020225
$\eta_2(1645)$	10225
$f_2(1810)$	9030225
$\eta_2(1870)$	10335
$f_2(1910)$	9040225
$f_2(1950)$	9050225
$f_2(2010)$	9060225
$f_2(2150)$	9070225
$f_2(2300)$	9080225
$f_2(2340)$	9090225
$\omega_3(1670)$	227
$\phi_3(1850)$	337
$f_4(2050)$	229
$f_J(2220)$	9000229
$f_4(2300)$	9010229

STRANGE MESONS		CHARMED MESONS		$c\bar{c}$ MESONS		LIGHT BARYONS		BOTTOM BARYONS	
$K_L^0$	130	$D^+$	411	$\eta_c(1S)$	441	$p$	2212	$\Lambda_b^0$	5122
$K_S^0$	310	$D^0$	421	$\chi_{c0}(1P)$	10441	$n$	2112	$\Sigma_b^-$	5112
$K^0$	311	$D_0^*(2400)^+$	10411	$\eta_c(2S)$	100441	$\Delta^{++}$	2224	$\Sigma_b^0$	5212
$K^+$	321	$D_0^*(2400)^0$	10421	$J/\psi(1S)$	443	$\Delta^+$	2214	$\Sigma_b^+$	5222
$K_0^*(800)^0$	9000311	$D^*(2010)^+$	413	$h_c(1P)$	10443	$\Delta^0$	2114	$\Sigma_b^{*-}$	5114
$K_0^*(800)^+$	9000321	$D^*(2007)^0$	423	$\chi_{c1}(1P)$	20443	$\Delta^-$	1114	$\Sigma_b^{*0}$	5214
$K_0^*(1430)^0$	10311	$D_1(2420)^+$	10413	$\psi(2S)$	100443	<b>STRANGE BARYONS</b>			
$K_0^*(1430)^+$	10321	$D_1(2420)^0$	10423	$\psi(3770)$	30443	$\Lambda$	3122	$\Sigma_b^{*+}$	5224
$K(1460)^0$	100311	$D_1(H)^+$	20413	$\psi(4040)$	9000443	$\Sigma^+$	3222	$\Xi_b^-$	5132
$K(1460)^+$	100321	$D_1(2430)^0$	20423	$\psi(4160)$	9010443	$\Sigma^0$	3212	$\Xi_b^0$	5232
$K(1830)^0$	9010311	$D_2^*(2460)^+$	415	$\psi(4415)$	9020443	$\Sigma^-$	3112	$\Xi_b'^-$	5312
$K(1830)^+$	9010321	$D_2^*(2460)^0$	425	$\chi_{c2}(1P)$	445	$\Sigma^{*+}$	3224 <sup>c</sup>	$\Xi_b'^0$	5322
$K_0^*(1950)^0$	9020311	$D_s^+$	431	$\chi_{c2}(2P)$	100445	$\Sigma^{*0}$	3214 <sup>c</sup>	$\Xi_b^{*-}$	5314
$K_0^*(1950)^+$	9020321	$D_{s0}^*(2317)^+$	10431	<b><math>b\bar{b}</math> MESONS</b>		$\Sigma^{*-}$	3114 <sup>c</sup>	$\Xi_b^{*0}$	5324
$K^*(892)^0$	313	$D_s^{*+}$	433	$\eta_b(1S)$	551	$\Xi^-$	3312	$\Omega_b^-$	5332
$K^*(892)^+$	323	$D_{s1}(2536)^+$	10433	$\chi_{b0}(1P)$	10551	$\Xi^{*0}$	3324 <sup>c</sup>	$\Omega_b^{*-}$	5334
$K_1(1270)^0$	10313	$D_{s1}(2460)^+$	20433	$\eta_b(2S)$	100551	$\Xi^{*-}$	3314 <sup>c</sup>	$\Xi_{bc}^0$	5142
$K_1(1270)^+$	10323	$D_{s2}^*(2573)^+$	435	$\chi_{b0}(2P)$	110551	$\Omega^-$	3334	$\Xi_{bc}^+$	5242
$K_1(1400)^0$	20313	<b>BOTTOM MESONS</b>		$\eta_b(3S)$	200551	<b>CHARMED BARYONS</b>			
$K_1(1400)^+$	20323	$B^0$	511	$\chi_{b0}(3P)$	210551	$\Lambda_c^+$	4122	$\Xi_{bc}'^0$	5412
$K^*(1410)^0$	100313	$B^+$	521	$\Upsilon(1S)$	553	$\Sigma_c^{++}$	4222	$\Xi_{bc}^{*+}$	5422
$K^*(1410)^+$	100323	$B_0^{*0}$	10511	$h_b(1P)$	10553	$\Sigma_c^+$	4212	$\Xi_{bc}^{*0}$	5414
$K_1(1650)^0$	9000313	$B_0^{*+}$	10521	$\chi_{b1}(1P)$	20553	$\Sigma_c^0$	4112	$\Xi_{bc}^{*+}$	5424
$K_1(1650)^+$	9000323	$B^{*0}$	513	$\Upsilon_1(1D)$	30553	$\Sigma_c^{*+}$	4224	$\Omega_{bc}^0$	5342
$K^*(1680)^0$	30313	$B^{*+}$	523	$\Upsilon(2S)$	100553	$\Sigma_c^{*0}$	4214	$\Omega_{bc}'^0$	5432
$K^*(1680)^+$	30323	$B_1(L)^0$	10513	$h_b(2P)$	110553	$\Sigma_c^+$	4114	$\Omega_{bc}^{*0}$	5434
$K_2^*(1430)^0$	315	$B_1(L)^+$	10523	$\chi_{b1}(2P)$	120553	$\Xi_c^+$	4232	$\Omega_{bcc}^+$	5442
$K_2^*(1430)^+$	325	$B_1(H)^0$	20513	$\Upsilon_1(2D)$	130553	$\Xi_c^0$	4132	$\Omega_{bcc}^{*+}$	5444
$K_2(1580)^0$	9000315	$B_1(H)^+$	20523	$\Upsilon(3S)$	200553	$\Xi_c'^+$	4322	$\Xi_{bb}^-$	5512
$K_2(1580)^+$	9000325	$B_2^0$	515	$h_b(3P)$	210553	$\Xi_c'^0$	4312	$\Xi_{bb}^0$	5522
$K_2(1770)^0$	10315	$B_2^{*+}$	525	$\chi_{b1}(3P)$	220553	$\Xi_c^{*+}$	4324	$\Xi_{bb}^{*-}$	5514
$K_2(1770)^+$	10325	$B_s^0$	531	$\Upsilon(4S)$	300553	$\Xi_c^{*0}$	4314	$\Xi_{bb}^{*0}$	5524
$K_2(1820)^0$	20315	$B_{s0}^{*0}$	10531	$\Upsilon(10860)$	9000553	$\Omega_c^0$	4332	$\Omega_{bb}^-$	5532
$K_2(1820)^+$	20325	$B_{s0}^{*+}$	533	$\Upsilon(11020)$	9010553	$\Omega_c^{*0}$	4334	$\Omega_{bb}^{*-}$	5534
$K_2^*(1980)^0$	9010315	$B_{s1}(L)^0$	10533	$\chi_{b2}(1P)$	555	$\Xi_{cc}^+$	4412	$\Omega_{bbc}^0$	5542
$K_2^*(1980)^+$	9010325	$B_{s1}(H)^0$	20533	$\eta_{b2}(1D)$	10555	$\Xi_{cc}^{*+}$	4422	$\Omega_{bbc}^{*0}$	5544
$K_2(2250)^0$	9020315	$B_{s2}^{*0}$	535	$\Upsilon_2(1D)$	20555	$\Xi_{cc}^{*+}$	4414	$\Omega_{bbb}^-$	5554
$K_2(2250)^+$	9020325	$B_c^+$	541	$\chi_{b2}(2P)$	100555	$\Xi_{cc}^{*+}$	4424		
$K_3^*(1780)^0$	317	$B_{c0}^{*+}$	10541	$\eta_{b2}(2D)$	110555	$\Omega_{cc}^+$	4432		
$K_3^*(1780)^+$	327	$B_c^{*+}$	543	$\Upsilon_2(2D)$	120555	$\Omega_{cc}^{*+}$	4434		
$K_3(2320)^0$	9010317	$B_{c1}(L)^+$	10543	$\chi_{b2}(3P)$	200555	$\Omega_{ccc}^{*+}$	4444		
$K_3(2320)^+$	9010327	$B_{c1}(H)^+$	20543	$\Upsilon_3(1D)$	557				
$K_4^*(2045)^0$	319	$B_{c2}^{*+}$	545	$\Upsilon_3(2D)$	100557				
$K_4^*(2045)^+$	329								
$K_4(2500)^0$	9000319								
$K_4(2500)^+$	9000329								

**Footnotes to the Tables:**

\*) Numbers or names in bold face are new or have changed since the 2014 *Review*.

a) Particular in the third generation, the left and right sfermion states may mix, as shown. The lighter mixed state is given the smaller number.

b) The physical  $\tilde{\chi}$  states are admixtures of the pure  $\tilde{\gamma}$ ,  $\tilde{Z}^0$ ,  $\tilde{W}^+$ ,  $\tilde{H}_1^0$ ,  $\tilde{H}_2^0$ , and  $\tilde{H}^+$  states.

c)  $\Sigma^*$  and  $\Xi^*$  are alternate names for  $\Sigma(1385)$  and  $\Xi(1530)$ .

# 44. CLEBSCH-GORDAN COEFFICIENTS, SPHERICAL HARMONICS, AND $d$ FUNCTIONS

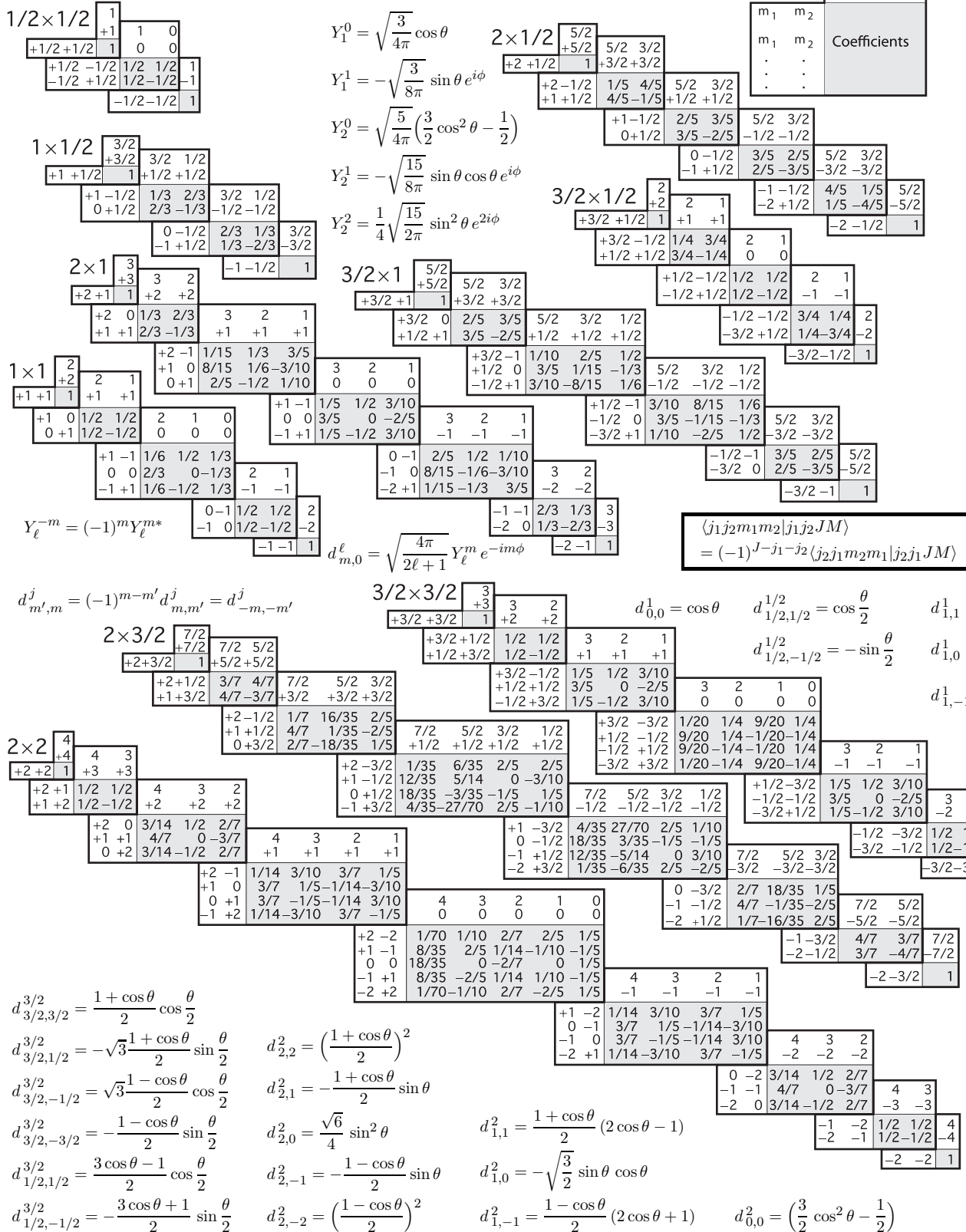
Note: A square-root sign is to be understood over *every* coefficient, e.g., for  $-8/15$  read  $-\sqrt{8/15}$ .

Notation:

$J$	$J$	...
$M$	$M$	...

$m_1$	$m_2$	
$m_1$	$m_2$	Coefficients
$\vdots$	$\vdots$	
$\vdots$	$\vdots$	



**Figure 44.1:** The sign convention is that of Wigner (*Group Theory*, Academic Press, New York, 1959), also used by Condon and Shortley (*The Theory of Atomic Spectra*, Cambridge Univ. Press, New York, 1953), Rose (*Elementary Theory of Angular Momentum*, Wiley, New York, 1957), and Cohen (*Tables of the Clebsch-Gordan Coefficients*, North American Rockwell Science Center, Thousand Oaks, Calif., 1974).

## 45. SU(3) ISOSCALAR FACTORS AND REPRESENTATION MATRICES

Written by R.L. Kelly (LBNL).

The most commonly used SU(3) isoscalar factors, corresponding to the singlet, octet, and decuplet content of  $8 \otimes 8$  and  $10 \otimes 8$ , are shown at the right. The notation uses particle names to identify the coefficients, so that the pattern of relative couplings may be seen at a glance. We illustrate the use of the coefficients below. See J.J. de Swart, Rev. Mod. Phys. **35**, 916 (1963) for detailed explanations and phase conventions.

A  $\sqrt{\phantom{x}}$  is to be understood over every integer in the matrices; the exponent 1/2 on each matrix is a reminder of this. For example, the  $\Xi \rightarrow \Omega K$  element of the  $10 \rightarrow 10 \otimes 8$  matrix is  $-\sqrt{6}/\sqrt{24} = -1/2$ .

Intramultiplet relative decay strengths may be read directly from the matrices. For example, in decuplet  $\rightarrow$  octet + octet decays, the ratio of  $\Omega^* \rightarrow \Xi \bar{K}$  and  $\Delta \rightarrow N \pi$  partial widths is, from the  $10 \rightarrow 8 \times 8$  matrix,

$$\frac{\Gamma(\Omega^* \rightarrow \Xi \bar{K})}{\Gamma(\Delta \rightarrow N \pi)} = \frac{12}{6} \times (\text{phase space factors}). \quad (45.1)$$

Including isospin Clebsch-Gordan coefficients, we obtain, *e.g.*,

$$\frac{\Gamma(\Omega^{*-} \rightarrow \Xi^0 K^-)}{\Gamma(\Delta^+ \rightarrow p \pi^0)} = \frac{1/2}{2/3} \times \frac{12}{6} \times p.s.f. = \frac{3}{2} \times p.s.f. \quad (45.2)$$

Partial widths for  $8 \rightarrow 8 \otimes 8$  involve a linear superposition of  $8_1$  (symmetric) and  $8_2$  (antisymmetric) couplings. For example,

$$\Gamma(\Xi^* \rightarrow \Xi \pi) \sim \left( -\sqrt{\frac{9}{20}} g_1 + \sqrt{\frac{3}{12}} g_2 \right)^2. \quad (45.3)$$

The relations between  $g_1$  and  $g_2$  (with de Swart's normalization) and the standard  $D$  and  $F$  couplings that appear in the interaction Lagrangian,

$$\mathcal{L} = -\sqrt{2} D \text{Tr}(\{\bar{B}, B\} M) + \sqrt{2} F \text{Tr}([\bar{B}, B] M), \quad (45.4)$$

where  $[\bar{B}, B] \equiv \bar{B}B - B\bar{B}$  and  $\{\bar{B}, B\} \equiv \bar{B}B + B\bar{B}$ , are

$$D = \frac{\sqrt{30}}{40} g_1, \quad F = \frac{\sqrt{6}}{24} g_2. \quad (45.5)$$

Thus, for example,

$$\Gamma(\Xi^* \rightarrow \Xi \pi) \sim (F - D)^2 \sim (1 - 2\alpha)^2, \quad (45.6)$$

where  $\alpha \equiv F/(D + F)$ . (This definition of  $\alpha$  is de Swart's. The alternative  $D/(D + F)$ , due to Gell-Mann, is also used.)

The generators of SU(3) transformations,  $\lambda_a$  ( $a = 1, 8$ ), are  $3 \times 3$  matrices that obey the following commutation and anticommutation relationships:

$$[\lambda_a, \lambda_b] \equiv \lambda_a \lambda_b - \lambda_b \lambda_a = 2i f_{abc} \lambda_c \quad (45.7)$$

$$\{\lambda_a, \lambda_b\} \equiv \lambda_a \lambda_b + \lambda_b \lambda_a = \frac{4}{3} \delta_{ab} I + 2d_{abc} \lambda_c, \quad (45.8)$$

where  $I$  is the  $3 \times 3$  identity matrix, and  $\delta_{ab}$  is the Kronecker delta symbol. The  $f_{abc}$  are odd under the permutation of any pair of indices, while the  $d_{abc}$  are even. The nonzero values are

$1 \rightarrow 8 \otimes 8$

$$(\Lambda) \rightarrow (N \bar{K} \Sigma \pi \Lambda \eta \Xi K) = \frac{1}{\sqrt{8}} \begin{pmatrix} 2 & 3 & -1 & -2 \end{pmatrix}^{1/2}$$

$8_1 \rightarrow 8 \otimes 8$

$$\begin{pmatrix} N \\ \Sigma \\ \Lambda \\ \Xi \end{pmatrix} \rightarrow \begin{pmatrix} N\pi & N\eta & \Sigma K & \Lambda K \\ N\bar{K} & \Sigma\pi & \Lambda\pi & \Sigma\eta & \Xi K \\ N\bar{K} & \Sigma\pi & \Lambda\eta & \Xi K \\ \Sigma\bar{K} & \Lambda\bar{K} & \Xi\pi & \Xi\eta \end{pmatrix} = \frac{1}{\sqrt{20}} \begin{pmatrix} 9 & -1 & -9 & -1 \\ -6 & 0 & 4 & 4 \\ 2 & -12 & -4 & -2 \\ 9 & -1 & -9 & -1 \end{pmatrix}^{1/2}$$

$8_2 \rightarrow 8 \otimes 8$

$$\begin{pmatrix} N \\ \Sigma \\ \Lambda \\ \Xi \end{pmatrix} \rightarrow \begin{pmatrix} N\pi & N\eta & \Sigma K & \Lambda K \\ N\bar{K} & \Sigma\pi & \Lambda\pi & \Sigma\eta & \Xi K \\ N\bar{K} & \Sigma\pi & \Lambda\eta & \Xi K \\ \Sigma\bar{K} & \Lambda\bar{K} & \Xi\pi & \Xi\eta \end{pmatrix} = \frac{1}{\sqrt{12}} \begin{pmatrix} 3 & 3 & 3 & -3 \\ 2 & 8 & 0 & 0 & -2 \\ 6 & 0 & 0 & 6 \\ 3 & 3 & 3 & -3 \end{pmatrix}^{1/2}$$

$10 \rightarrow 8 \otimes 8$

$$\begin{pmatrix} \Delta \\ \Sigma \\ \Xi \\ \Omega \end{pmatrix} \rightarrow \begin{pmatrix} N\pi & \Sigma K \\ N\bar{K} & \Sigma\pi & \Lambda\pi & \Sigma\eta & \Xi K \\ \Sigma\bar{K} & \Lambda\bar{K} & \Xi\pi & \Xi\eta \\ \Xi\bar{K} \end{pmatrix} = \frac{1}{\sqrt{12}} \begin{pmatrix} -6 & 6 \\ -2 & 2 & -3 & 3 & 2 \\ 3 & -3 & 3 & 3 \\ 12 \end{pmatrix}^{1/2}$$

$8 \rightarrow 10 \otimes 8$

$$\begin{pmatrix} N \\ \Sigma \\ \Lambda \\ \Xi \end{pmatrix} \rightarrow \begin{pmatrix} \Delta\pi & \Sigma K \\ \Delta\bar{K} & \Sigma\pi & \Sigma\eta & \Xi K \\ \Sigma\pi & \Xi K \\ \Sigma\bar{K} & \Xi\pi & \Xi\eta & \Omega K \end{pmatrix} = \frac{1}{\sqrt{15}} \begin{pmatrix} -12 & 3 \\ 8 & -2 & -3 & 2 \\ -9 & 6 \\ 3 & -3 & -3 & 6 \end{pmatrix}^{1/2}$$

$10 \rightarrow 10 \otimes 8$

$$\begin{pmatrix} \Delta \\ \Sigma \\ \Xi \\ \Omega \end{pmatrix} \rightarrow \begin{pmatrix} \Delta\pi & \Delta\eta & \Sigma K \\ \Delta\bar{K} & \Sigma\pi & \Sigma\eta & \Xi K \\ \Sigma\bar{K} & \Xi\pi & \Xi\eta & \Omega K \\ \Xi\bar{K} & \Omega\eta \end{pmatrix} = \frac{1}{\sqrt{24}} \begin{pmatrix} 15 & 3 & -6 \\ 8 & 8 & 0 & -8 \\ 12 & 3 & -3 & -6 \\ 12 & -12 \end{pmatrix}^{1/2}$$

$abc$	$f_{abc}$	$abc$	$d_{abc}$	$abc$	$d_{abc}$
123	1	118	$1/\sqrt{3}$	355	1/2
147	1/2	146	1/2	366	-1/2
156	-1/2	157	1/2	377	-1/2
246	1/2	228	$1/\sqrt{3}$	448	$-1/(2\sqrt{3})$
257	1/2	247	-1/2	558	$-1/(2\sqrt{3})$
345	1/2	256	1/2	668	$-1/(2\sqrt{3})$
367	-1/2	338	$1/\sqrt{3}$	778	$-1/(2\sqrt{3})$
458	$\sqrt{3}/2$	344	1/2	888	$-1/\sqrt{3}$
678	$\sqrt{3}/2$				

The  $\lambda_a$ 's are

$$\lambda_1 = \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad \lambda_2 = \begin{pmatrix} 0 & -i & 0 \\ i & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad \lambda_3 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$\lambda_4 = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix} \quad \lambda_5 = \begin{pmatrix} 0 & 0 & -i \\ 0 & 0 & 0 \\ i & 0 & 0 \end{pmatrix} \quad \lambda_6 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$$

$$\lambda_7 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & -i \\ 0 & i & 0 \end{pmatrix} \quad \lambda_8 = \frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -2 \end{pmatrix}$$

Equation (45.7) defines the Lie algebra of SU(3). A general  $d$ -dimensional representation is given by a set of  $d \times d$  matrices satisfying Eq. (45.7) with the  $f_{abc}$  given above. Equation (45.8) is specific to the defining 3-dimensional representation.

## 46. $SU(n)$ MULTIPLETS AND YOUNG DIAGRAMS

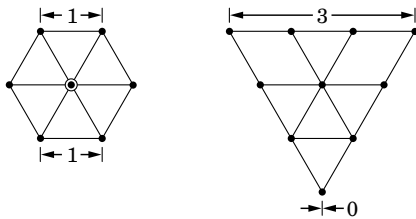
Written by C.G. Wohl (LBNL).

This note tells (1) how  $SU(n)$  particle multiplets are identified or labeled, (2) how to find the number of particles in a multiplet from its label, (3) how to draw the Young diagram for a multiplet, and (4) how to use Young diagrams to determine the overall multiplet structure of a composite system, such as a 3-quark or a meson-baryon system.

In much of the literature, the word “representation” is used where we use “multiplet,” and “tableau” is used where we use “diagram.”

### 46.1. Multiplet labels

An  $SU(n)$  multiplet is uniquely identified by a string of  $(n-1)$  nonnegative integers:  $(\alpha, \beta, \gamma, \dots)$ . Any such set of integers specifies a multiplet. For an  $SU(2)$  multiplet such as an isospin multiplet, the single integer  $\alpha$  is the number of *steps* from one end of the multiplet to the other (*i.e.*, it is one fewer than the number of particles in the multiplet). In  $SU(3)$ , the two integers  $\alpha$  and  $\beta$  are the numbers of steps across the top and bottom levels of the multiplet diagram. Thus the labels for the  $SU(3)$  octet and decuplet



are (1,1) and (3,0). For larger  $n$ , the interpretation of the integers in terms of the geometry of the multiplets, which exist in an  $(n-1)$ -dimensional space, is not so readily apparent.

The label for the  $SU(n)$  singlet is  $(0, 0, \dots, 0)$ . In a flavor  $SU(n)$ , the  $n$  quarks together form a  $(1, 0, \dots, 0)$  multiplet, and the  $n$  antiquarks belong to a  $(0, \dots, 0, 1)$  multiplet. These two multiplets are *conjugate* to one another, which means their labels are related by  $(\alpha, \beta, \dots) \leftrightarrow (\dots, \beta, \alpha)$ .

## 46.2. Number of particles

The number of particles in a multiplet,  $N = N(\alpha, \beta, \dots)$ , is given as follows (note the pattern of the equations).

In  $SU(2)$ ,  $N = N(\alpha)$  is

$$N = \frac{(\alpha + 1)}{1} . \quad (46.1)$$

In  $SU(3)$ ,  $N = N(\alpha, \beta)$  is

$$N = \frac{(\alpha+1)}{1} \cdot \frac{(\beta+1)}{1} \cdot \frac{(\alpha+\beta+2)}{2} . \quad (46.2)$$

In  $SU(4)$ ,  $N = N(\alpha, \beta, \gamma)$  is

$$N = \frac{(\alpha+1)}{1} \cdot \frac{(\beta+1)}{1} \cdot \frac{(\gamma+1)}{1} \cdot \frac{(\alpha+\beta+2)}{2} \cdot \frac{(\beta+\gamma+2)}{2} \cdot \frac{(\alpha+\beta+\gamma+3)}{3} \quad (46.3)$$

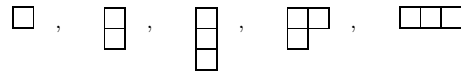
Note that in Eq. (46.3) there is no factor with  $(\alpha + \gamma + 2)$ : only a *consecutive* sequence of the label integers appears in any factor. One more example should make the pattern clear for any  $SU(n)$ . In  $SU(5)$ ,  $N = N(\alpha, \beta, \gamma, \delta)$  is

$$N = \frac{(\alpha+1)}{1} \cdot \frac{(\beta+1)}{1} \cdot \frac{(\gamma+1)}{1} \cdot \frac{(\delta+1)}{1} \cdot \frac{(\alpha+\beta+2)}{2} \cdot \frac{(\beta+\gamma+2)}{2} \\ \times \frac{(\gamma+\delta+2)}{2} \cdot \frac{(\alpha+\beta+\gamma+3)}{3} \cdot \frac{(\beta+\gamma+\delta+3)}{3} \cdot \frac{(\alpha+\beta+\gamma+\delta+4)}{4} \quad (46.4)$$

From the symmetry of these equations, it is clear that multiplets that are conjugate to one another have the same number of particles, but so can other multiplets. For example, the  $SU(4)$  multiplets  $(3,0,0)$  and  $(1,1,0)$  each have 20 particles. Try the equations and see.

### 46.3. Young diagrams

A Young diagram consists of an array of boxes (or some other symbol) arranged in one or more *left-justified* rows, with each row being *at least as long* as the row beneath. The correspondence between a diagram and a multiplet label is: The top row juts out  $\alpha$  boxes to the right past the end of the second row, the second row juts out  $\beta$  boxes to the right past the end of the third row, *etc.* A diagram in  $SU(n)$  has at most  $n$  rows. There can be any number of “completed” columns of  $n$  boxes buttressing the left of a diagram; these don’t affect the label. Thus in  $SU(3)$  the diagrams





represent the multiplets (1,0), (0,1), (0,0), (1,1), and (3,0). In any  $SU(n)$ , the quark multiplet is represented by a single box, the antiquark multiplet by a column of  $(n-1)$  boxes, and a singlet by a completed column of  $n$  boxes.

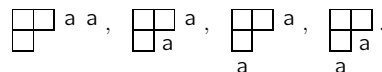
#### 46.4. Coupling multiplets together

The following recipe tells how to find the multiplets that occur in coupling two multiplets together. To couple together more than two multiplets, first couple two, then couple a third with each of the multiplets obtained from the first two, *etc.*

First a definition: A sequence of the letters  $a, b, c, \dots$  is *admissible* if at any point in the sequence at least as many  $a$ 's have occurred as  $b$ 's, at least as many  $b$ 's have occurred as  $c$ 's, *etc.* Thus  $abcd$  and  $aabcb$  are admissible sequences and  $abb$  and  $acb$  are not. Now the recipe:

(a) Draw the Young diagrams for the two multiplets, but in one of the diagrams replace the boxes in the first row with  $a$ 's, the boxes in the second row with  $b$ 's, *etc.* Thus, to couple two SU(3) octets (such as the  $\pi$ -meson octet and the baryon octet), we start with  and . The *unlettered* diagram forms the *upper left-hand corner* of all the enlarged diagrams constructed below.

(b) Add the  $a$ 's from the lettered diagram to the right-hand ends of the rows of the unlettered diagram to form all possible legitimate Young diagrams that have no more than one  $a$  per column. In general, there will be several distinct diagrams, and all the  $a$ 's appear in each diagram. At this stage, for the coupling of the two  $SU(3)$  octets, we have:



(c) Use the  $b$ 's to further enlarge the diagrams already obtained, subject to the same rules. Then throw away any diagram in which the full sequence of letters formed by reading *right to left* in the first row, then the second row, *etc.*, is not admissible.

(d) Proceed as in (c) with the  $c$ 's (if any), *etc.*

The final result of the coupling of the two SU(3) octets is:



$$\begin{array}{c} \begin{array}{|c|c|} \hline & \\ \hline \end{array} a & a \oplus \begin{array}{|c|c|} \hline & \\ \hline \end{array} a & a \oplus \begin{array}{|c|c|} \hline & \\ \hline \end{array} a & \begin{array}{|c|c|} \hline & \\ \hline \end{array} a \oplus \begin{array}{|c|c|} \hline & \\ \hline \end{array} a & \begin{array}{|c|c|} \hline & \\ \hline \end{array} a & \begin{array}{|c|c|} \hline & \\ \hline \end{array} a \\ \begin{array}{|c|c|} \hline & \\ \hline \end{array} b & & & & & \\ & b & & b & a & a & b \end{array}$$

Here only the diagrams with admissible sequences of  $a$ 's and  $b$ 's and with fewer than four rows (since  $n = 3$ ) have been kept. In terms of multiplet labels, the above may be written

$$(1,1) \otimes (1,1) = (2,2) \oplus (3,0) \oplus (0,3) \oplus (1,1) \oplus (1,1) \oplus (0,0) \text{ .}$$

In terms of numbers of particles, it may be written

$$8 \otimes 8 = 27 \oplus 10 \oplus \overline{10} \oplus 8 \oplus 8 \oplus 1.$$

The product of the numbers on the left here is equal to the sum on the right, a useful check. (See also Sec. 15 on the Quark Model.)

## 47. KINEMATICS

Revised January 2000 by J.D. Jackson (LBNL) and January 2016 by D.R. Tovey (Sheffield).

Throughout this section units are used in which  $\hbar = c = 1$ . The following conversions are useful:  $\hbar c = 197.3 \text{ MeV fm}$ ,  $(\hbar c)^2 = 0.3894 \text{ (GeV)}^2 \text{ mb}$ .

## 47.1. Lorentz transformations

The energy  $E$  and 3-momentum  $\mathbf{p}$  of a particle of mass  $m$  form a 4-vector  $p = (E, \mathbf{p})$  whose square  $p^2 \equiv E^2 - |\mathbf{p}|^2 = m^2$ . The velocity of the particle is  $\boldsymbol{\beta} = \mathbf{p}/E$ . The energy and momentum  $(E^*, \mathbf{p}^*)$  viewed from a frame moving with velocity  $\boldsymbol{\beta}_f$  are given by

$$\begin{pmatrix} E^* \\ \mathbf{p}^* \end{pmatrix} = \begin{pmatrix} \gamma_f & -\gamma_f \boldsymbol{\beta}_f \\ -\gamma_f \boldsymbol{\beta}_f & \gamma_f \end{pmatrix} \begin{pmatrix} E \\ \mathbf{p} \end{pmatrix}, \quad p_T^* = p_T, \quad (47.1)$$

where  $\gamma_f = (1 - \beta_f^2)^{-1/2}$  and  $p_T$  ( $p_{\parallel}$ ) are the components of  $\mathbf{p}$  perpendicular (parallel) to  $\boldsymbol{\beta}_f$ . Other 4-vectors, such as the space-time coordinates of events, of course transform in the same way. The scalar product of two 4-momenta  $\mathbf{p}_1 \cdot \mathbf{p}_2 = E_1 E_2 - \mathbf{p}_1 \cdot \mathbf{p}_2$  is invariant (frame independent).

## 47.2. Center-of-mass energy and momentum

In the collision of two particles of masses  $m_1$  and  $m_2$  the total center-of-mass energy can be expressed in the Lorentz-invariant form

$$\begin{aligned} E_{\text{cm}} &= \left[ (E_1 + E_2)^2 - (\mathbf{p}_1 + \mathbf{p}_2)^2 \right]^{1/2}, \\ &= \left[ m_1^2 + m_2^2 + 2E_1 E_2 (1 - \beta_1 \beta_2 \cos \theta) \right]^{1/2}, \end{aligned} \quad (47.2)$$

where  $\theta$  is the angle between the particles. In the frame where one particle (of mass  $m_2$ ) is at rest (lab frame),

$$E_{\text{cm}} = (m_1^2 + m_2^2 + 2E_{1\text{lab}} m_2)^{1/2}. \quad (47.3)$$

The velocity of the center-of-mass in the lab frame is

$$\boldsymbol{\beta}_{\text{cm}} = \mathbf{p}_{\text{lab}} / (E_{1\text{lab}} + m_2), \quad (47.4)$$

where  $\mathbf{p}_{\text{lab}} \equiv \mathbf{p}_{1\text{lab}}$  and

$$\gamma_{\text{cm}} = (E_{1\text{lab}} + m_2) / E_{\text{cm}}. \quad (47.5)$$

The c.m. momenta of particles 1 and 2 are of magnitude

$$p_{\text{cm}} = p_{\text{lab}} \frac{m_2}{E_{\text{cm}}}. \quad (47.6)$$

For example, if a 0.80 GeV/c kaon beam is incident on a proton target, the center of mass energy is 1.699 GeV and the center of mass momentum of either particle is 0.442 GeV/c. It is also useful to note that

$$E_{\text{cm}} dE_{\text{cm}} = m_2 dE_{1\text{lab}} = m_2 \beta_{1\text{lab}} dp_{\text{lab}}. \quad (47.7)$$

## 47.3. Lorentz-invariant amplitudes

The matrix elements for a scattering or decay process are written in terms of an invariant amplitude  $-i\mathcal{M}$ . As an example, the  $S$ -matrix for  $2 \rightarrow 2$  scattering is related to  $\mathcal{M}$  by

$$\begin{aligned} \langle p'_1 p'_2 | S | p_1 p_2 \rangle &= I - i(2\pi)^4 \delta^4(p_1 + p_2 - p'_1 - p'_2) \\ &\times \frac{\mathcal{M}(p_1, p_2; p'_1, p'_2)}{(2E_1)^{1/2} (2E_2)^{1/2} (2E'_1)^{1/2} (2E'_2)^{1/2}}. \end{aligned} \quad (47.8)$$

The state normalization is such that

$$\langle p' | p \rangle = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}'). \quad (47.9)$$

For a  $2 \rightarrow 2$  scattering process producing unstable particles  $1'$  and  $2'$  decaying via  $1' \rightarrow 3'4'$  and  $2' \rightarrow 5'6'$  the matrix element for the complete process can be written in the narrow width approximation as:

$$\begin{aligned} &\mathcal{M}(12 \rightarrow 3'4'5'6') = \\ &\sum_{h_1, h_2} \frac{\mathcal{M}(12 \rightarrow 1'2') \mathcal{M}(1' \rightarrow 3'4') \mathcal{M}(2' \rightarrow 5'6')}{(m_{3'4'}^2 - m_{1'}^2 + im_{1'}\Gamma_{1'})(m_{5'6'}^2 - m_{2'}^2 + im_{2'}\Gamma_{2'})}. \end{aligned} \quad (47.10)$$

Here,  $m_{ij}$  is the invariant mass of particles  $i$  and  $j$ ,  $m_k$  and  $\Gamma_k$  are the mass and total width of particle  $k$ , and the sum runs over the helicities of the intermediate particles. This enables the cross section for such a process to be written as the product of the cross section for the initial  $2 \rightarrow 2$  scattering process with the branching ratios (relative partial decay rates) of the subsequent decays.

## 47.4. Particle decays

The partial decay rate of a particle of mass  $M$  into  $n$  bodies in its rest frame is given in terms of the Lorentz-invariant matrix element  $\mathcal{M}$  by

$$d\Gamma = \frac{(2\pi)^4}{2M} |\mathcal{M}|^2 d\Phi_n(P; p_1, \dots, p_n), \quad (47.11)$$

where  $d\Phi_n$  is an element of  $n$ -body phase space given by

$$d\Phi_n(P; p_1, \dots, p_n) = \delta^4(P - \sum_{i=1}^n p_i) \prod_{i=1}^n \frac{d^3 p_i}{(2\pi)^3 2E_i}. \quad (47.12)$$

This phase space can be generated recursively, viz.

$$\begin{aligned} d\Phi_n(P; p_1, \dots, p_n) &= d\Phi_j(q; p_1, \dots, p_j) \\ &\times d\Phi_{n-j+1}(P; q, p_{j+1}, \dots, p_n) (2\pi)^3 dq^2, \end{aligned} \quad (47.13)$$

where  $q^2 = (\sum_{i=1}^j E_i)^2 - |\sum_{i=1}^j \mathbf{p}_i|^2$ . This form is particularly useful in the case where a particle decays into another particle that subsequently decays.

**47.4.1. Survival probability:** If a particle of mass  $M$  has mean proper lifetime  $\tau$  ( $= 1/\Gamma$ ) and has momentum  $(E, \mathbf{p})$ , then the probability that it lives for a time  $t_0$  or greater before decaying is given by

$$P(t_0) = e^{-t_0 \Gamma / \gamma} = e^{-Mt_0 \Gamma / E}, \quad (47.14)$$

and the probability that it travels a distance  $x_0$  or greater is

$$P(x_0) = e^{-Mx_0 \Gamma / |\mathbf{p}|}. \quad (47.15)$$

## 47.4.2. Two-body decays :

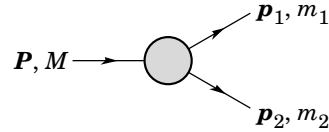


Figure 47.1: Definitions of variables for two-body decays.

In the rest frame of a particle of mass  $M$ , decaying into 2 particles labeled 1 and 2,

$$E_1 = \frac{M^2 - m_2^2 + m_1^2}{2M}, \quad (47.16)$$

$$|\mathbf{p}_1| = |\mathbf{p}_2|$$

$$= \frac{[(M^2 - (m_1 + m_2)^2)(M^2 - (m_1 - m_2)^2)]^{1/2}}{2M}, \quad (47.17)$$

and

$$d\Gamma = \frac{1}{32\pi^2} |\mathcal{M}|^2 \frac{|\mathbf{p}_1|}{M^2} d\Omega, \quad (47.18)$$

where  $d\Omega = d\phi_1 d(\cos \theta_1)$  is the solid angle of particle 1. The invariant mass  $M$  can be determined from the energies and momenta using Eq. (47.2) with  $M = E_{\text{cm}}$ .

## 47.4.3. Three-body decays :

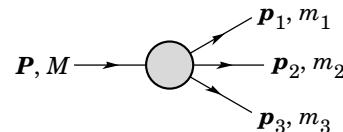


Figure 47.2: Definitions of variables for three-body decays.



Defining  $p_{ij} = p_i + p_j$  and  $m_{ij}^2 = p_{ij}^2$ , then  $m_{12}^2 + m_{23}^2 + m_{13}^2 = M^2 + m_1^2 + m_2^2 + m_3^2$  and  $m_{12}^2 = (P - p_3)^2 = M^2 + m_3^2 - 2ME_3$ , where  $E_3$  is the energy of particle 3 in the rest frame of  $M$ . In that frame, the momenta of the three decay particles lie in a plane. The relative orientation of these three momenta is fixed if their energies are known. The momenta can therefore be specified in space by giving three Euler angles  $(\alpha, \beta, \gamma)$  that specify the orientation of the final system relative to the initial particle [1]. Then

$$d\Gamma = \frac{1}{(2\pi)^5} \frac{1}{16M} |\mathcal{M}|^2 dE_1 dE_3 d\alpha d(\cos\beta) d\gamma. \quad (47.19)$$

Alternatively

$$d\Gamma = \frac{1}{(2\pi)^5} \frac{1}{16M^2} |\mathcal{M}|^2 |\mathbf{p}_1^*| |\mathbf{p}_3| dm_{12} d\Omega_1^* d\Omega_3, \quad (47.20)$$

where  $(|\mathbf{p}_1^*|, \Omega_1^*)$  is the momentum of particle 1 in the rest frame of 1 and 2, and  $\Omega_3$  is the angle of particle 3 in the rest frame of the decaying particle.  $|\mathbf{p}_1^*|$  and  $|\mathbf{p}_3|$  are given by

$$|\mathbf{p}_1^*| = \frac{[(m_{12}^2 - (m_1 + m_2)^2)(m_{12}^2 - (m_1 - m_2)^2)]^{1/2}}{2m_{12}}, \quad (47.21a)$$

and

$$|\mathbf{p}_3| = \frac{[(M^2 - (m_{12} + m_3)^2)(M^2 - (m_{12} - m_3)^2)]^{1/2}}{2M}. \quad (47.21b)$$

[Compare with Eq. (47.17).]

If the decaying particle is a scalar or we average over its spin states, then integration over the angles in Eq. (47.19) gives

$$\begin{aligned} d\Gamma &= \frac{1}{(2\pi)^3} \frac{1}{8M} |\mathcal{M}|^2 dE_1 dE_3 \\ &= \frac{1}{(2\pi)^3} \frac{1}{32M^3} |\mathcal{M}|^2 dm_{12}^2 dm_{23}^2. \end{aligned} \quad (47.22)$$

This is the standard form for the Dalitz plot.

**47.4.3.1. Dalitz plot:** For a given value of  $m_{12}^2$ , the range of  $m_{23}^2$  is determined by its values when  $\mathbf{p}_2$  is parallel or antiparallel to  $\mathbf{p}_3$ :

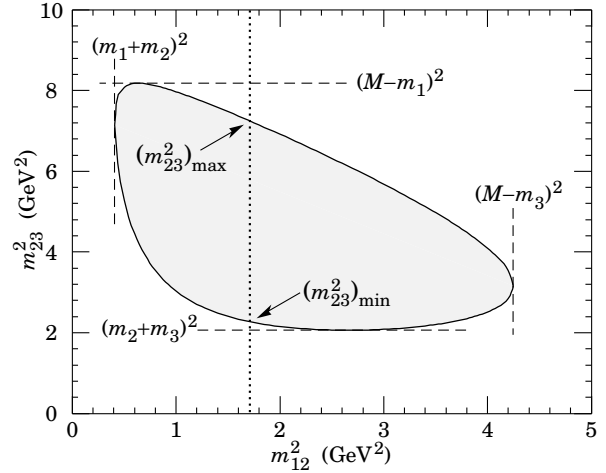
$$(m_{23}^2)_{\max} = (E_2^* + E_3^*)^2 - \left( \sqrt{E_2^{*2} - m_2^2} - \sqrt{E_3^{*2} - m_3^2} \right)^2, \quad (47.23a)$$

$$(m_{23}^2)_{\min} = (E_2^* + E_3^*)^2 - \left( \sqrt{E_2^{*2} - m_2^2} + \sqrt{E_3^{*2} - m_3^2} \right)^2. \quad (47.23b)$$

Here  $E_2^* = (m_{12}^2 - m_1^2 + m_2^2)/2m_{12}$  and  $E_3^* = (M^2 - m_{12}^2 - m_3^2)/2m_{12}$  are the energies of particles 2 and 3 in the  $m_{12}$  rest frame. The scatter plot in  $m_{12}^2$  and  $m_{23}^2$  is called a Dalitz plot. If  $|\mathcal{M}|^2$  is constant, the allowed region of the plot will be uniformly populated with events [see Eq. (47.22)]. A nonuniformity in the plot gives immediate information on  $|\mathcal{M}|^2$ . For example, in the case of  $D \rightarrow K\pi\pi$ , bands appear when  $m_{(K\pi)} = m_{K^*(892)}$ , reflecting the appearance of the decay chain  $D \rightarrow K^*(892)\pi \rightarrow K\pi\pi$ .

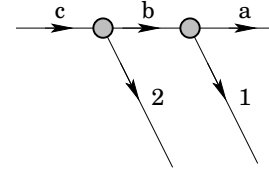
#### 47.4.4. Kinematic limits :

**47.4.4.1. Three-body decays:** In a three-body decay (Fig. 47.2) the maximum of  $|\mathbf{p}_3|$ , [given by Eq. (47.21)], is achieved when  $m_{12} = m_1 + m_2$ , i.e., particles 1 and 2 have the same vector velocity in the rest frame of the decaying particle. If, in addition,  $m_3 > m_1, m_2$ , then  $|\mathbf{p}_3|_{\max} > |\mathbf{p}_1|_{\max}, |\mathbf{p}_2|_{\max}$ . The distribution of  $m_{12}$  values possesses an end-point or maximum value at  $m_{12} = M - m_3$ . This can be used to constrain the mass difference of a parent particle and one invisible decay product.



**Figure 47.3:** Dalitz plot for a three-body final state. In this example, the state is  $\pi^+\bar{K}^0 p$  at 3 GeV. Four-momentum conservation restricts events to the shaded region.

#### 47.4.4.2. Sequential two-body decays:



**Figure 47.4:** Particles participating in sequential two-body decay chain. Particles labeled 1 and 2 are visible while the particle terminating the chain (a) is invisible.

When a heavy particle initiates a sequential chain of two-body decays terminating in an invisible particle, constraints on the masses of the states participating in the chain can be obtained from end-points and thresholds in invariant mass distributions of the aggregated decay products. For the two-step decay chain depicted in Fig. 47.4 the invariant mass distribution of the two visible particles possesses an end-point given by:

$$(m_{12}^{\max})^2 = \frac{(m_c^2 - m_b^2)(m_b^2 - m_a^2)}{m_b^2}, \quad (47.24)$$

provided particles 1 and 2 are massless. If visible particle 1 has non-zero mass  $m_1$  then Eq. (47.24) is replaced by

$$(m_{12}^{\max})^2 = m_1^2 + \frac{(m_c^2 - m_b^2)}{2m_b^2} \times \left( m_1^2 + m_b^2 - m_a^2 + \sqrt{(-m_1^2 + m_b^2 - m_a^2)^2 - 4m_1^2 m_a^2} \right). \quad (47.25)$$

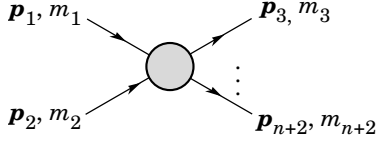
See Refs. 2 and 3 for other cases.

**47.4.4.5. Multibody decays :** The above results may be generalized to final states containing any number of particles by combining some of the particles into “effective particles” and treating the final states as 2 or 3 “effective particle” states. Thus, if  $p_{ijk\dots} = p_i + p_j + p_k + \dots$ , then

$$m_{ijk\dots} = \sqrt{p_{ijk\dots}^2}, \quad (47.26)$$

and  $m_{ijk\dots}$  may be used in place of e.g.,  $m_{12}$  in the relations in Sec. 47.4.3 or Sec. 47.4.4 above.

## 47.5. Cross sections



**Figure 47.5:** Definitions of variables for production of an  $n$ -body final state.

The differential cross section is given by

$$d\sigma = \frac{(2\pi)^4 |\mathcal{M}|^2}{4\sqrt{(p_1 \cdot p_2)^2 - m_1^2 m_2^2}} \times d\Phi_n(p_1 + p_2; p_3, \dots, p_{n+2}) . \quad (47.27)$$

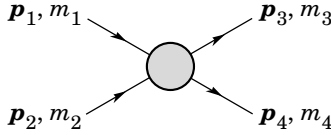
[See Eq. (47.12).] In the rest frame of  $m_2(\text{lab})$ ,

$$\sqrt{(p_1 \cdot p_2)^2 - m_1^2 m_2^2} = m_2 p_{1\text{lab}} ; \quad (47.28a)$$

while in the center-of-mass frame

$$\sqrt{(p_1 \cdot p_2)^2 - m_1^2 m_2^2} = p_{1\text{cm}} \sqrt{s} . \quad (47.28b)$$

## 47.5.1. Two-body reactions :



**Figure 47.6:** Definitions of variables for a two-body final state.

Two particles of momenta  $p_1$  and  $p_2$  and masses  $m_1$  and  $m_2$  scatter to particles of momenta  $p_3$  and  $p_4$  and masses  $m_3$  and  $m_4$ ; the Lorentz-invariant Mandelstam variables are defined by

$$s = (p_1 + p_2)^2 = (p_3 + p_4)^2 = m_1^2 + 2E_1 E_2 - 2\mathbf{p}_1 \cdot \mathbf{p}_2 + m_2^2 , \quad (47.29)$$

$$t = (p_1 - p_3)^2 = (p_2 - p_4)^2 = m_1^2 - 2E_1 E_3 + 2\mathbf{p}_1 \cdot \mathbf{p}_3 + m_3^2 , \quad (47.30)$$

$$u = (p_1 - p_4)^2 = (p_2 - p_3)^2 = m_1^2 - 2E_1 E_4 + 2\mathbf{p}_1 \cdot \mathbf{p}_4 + m_4^2 , \quad (47.31)$$

and they satisfy

$$s + t + u = m_1^2 + m_2^2 + m_3^2 + m_4^2 . \quad (47.32)$$

The two-body cross section may be written as

$$\frac{d\sigma}{dt} = \frac{1}{64\pi s} \frac{1}{|\mathbf{p}_{1\text{cm}}|^2} |\mathcal{M}|^2 . \quad (47.33)$$

In the center-of-mass frame

$$t = (E_{1\text{cm}} - E_{3\text{cm}})^2 - (p_{1\text{cm}} - p_{3\text{cm}})^2 - 4p_{1\text{cm}} p_{3\text{cm}} \sin^2(\theta_{\text{cm}}/2) = t_0 - 4p_{1\text{cm}} p_{3\text{cm}} \sin^2(\theta_{\text{cm}}/2) , \quad (47.34)$$

where  $\theta_{\text{cm}}$  is the angle between particle 1 and 3. The limiting values  $t_0$  ( $\theta_{\text{cm}} = 0$ ) and  $t_1$  ( $\theta_{\text{cm}} = \pi$ ) for  $2 \rightarrow 2$  scattering are

$$t_0(t_1) = \left[ \frac{m_1^2 - m_3^2 - m_2^2 + m_4^2}{2\sqrt{s}} \right]^2 - (p_{1\text{cm}} \mp p_{3\text{cm}})^2 . \quad (47.35)$$

In the literature the notation  $t_{\min}$  ( $t_{\max}$ ) for  $t_0$  ( $t_1$ ) is sometimes used, which should be discouraged since  $t_0 > t_1$ . The center-of-mass energies and momenta of the incoming particles are

$$E_{1\text{cm}} = \frac{s + m_1^2 - m_2^2}{2\sqrt{s}} , \quad E_{2\text{cm}} = \frac{s + m_2^2 - m_1^2}{2\sqrt{s}} , \quad (47.36)$$

For  $E_{3\text{cm}}$  and  $E_{4\text{cm}}$ , change  $m_1$  to  $m_3$  and  $m_2$  to  $m_4$ . Then

$$p_{i\text{cm}} = \sqrt{E_{i\text{cm}}^2 - m_i^2} \text{ and } p_{1\text{cm}} = \frac{p_{1\text{lab}} m_2}{\sqrt{s}} . \quad (47.37)$$

Here the subscript lab refers to the frame where particle 2 is at rest. [For other relations see Eqs. (47.2)–(47.4).]

**47.5.2. Inclusive reactions :** Choose some direction (usually the beam direction) for the  $z$ -axis; then the energy and momentum of a particle can be written as

$$E = m_T \cosh y , \quad p_x , p_y , p_z = m_T \sinh y , \quad (47.38)$$

where  $m_T$ , conventionally called the ‘transverse mass’, is given by

$$m_T^2 = m^2 + p_x^2 + p_y^2 . \quad (47.39)$$

and the rapidity  $y$  is defined by

$$y = \frac{1}{2} \ln \left( \frac{E + p_z}{E - p_z} \right)$$

$$= \ln \left( \frac{E + p_z}{m_T} \right) = \tanh^{-1} \left( \frac{p_z}{E} \right) . \quad (47.40)$$

Note that the definition of the transverse mass in Eq. (47.39) differs from that used by experimentalists at hadron colliders (see Sec. 47.6.1 below). Under a boost in the  $z$ -direction to a frame with velocity  $\beta$ ,  $y \rightarrow y - \tanh^{-1} \beta$ . Hence the shape of the rapidity distribution  $dN/dy$  is invariant, as are differences in rapidity. The invariant cross section may also be rewritten

$$E \frac{d^3\sigma}{d^3p} = \frac{d^3\sigma}{d\phi dy p_T dp_T} \Rightarrow \frac{d^2\sigma}{\pi dy d(p_T^2)} . \quad (47.41)$$

The second form is obtained using the identity  $dy/dp_z = 1/E$ , and the third form represents the average over  $\phi$ .

Feynman’s  $x$  variable is given by

$$x = \frac{p_z}{p_{z\text{max}}} \approx \frac{E + p_z}{(E + p_z)_{\text{max}}} \quad (p_T \ll |p_z|) . \quad (47.42)$$

In the c.m. frame,

$$x \approx \frac{2p_{z\text{cm}}}{\sqrt{s}} = \frac{2m_T \sinh y_{\text{cm}}}{\sqrt{s}} \quad (47.43)$$

and

$$= (y_{\text{cm}})_{\text{max}} = \ln(\sqrt{s}/m) . \quad (47.44)$$

The invariant mass  $M$  of the two-particle system described in Sec. 47.4.2 can be written in terms of these variables as

$$M^2 = m_1^2 + m_2^2 + 2[E_T(1)E_T(2) \cosh \Delta y - \mathbf{p}_T(1) \cdot \mathbf{p}_T(2)] , \quad (47.45)$$

where

$$E_T(i) = \sqrt{|\mathbf{p}_T(i)|^2 + m_i^2} , \quad (47.46)$$

and  $\mathbf{p}_T(i)$  denotes the transverse momentum vector of particle  $i$ .

For  $p \gg m$ , the rapidity [Eq. (47.40)] may be expanded to obtain

$$y = \frac{1}{2} \ln \frac{\cos^2(\theta/2) + m^2/4p^2 + \dots}{\sin^2(\theta/2) + m^2/4p^2 + \dots} \approx -\ln \tan(\theta/2) \equiv \eta \quad (47.47)$$

where  $\cos \theta = p_z/p$ . The pseudorapidity  $\eta$  defined by the second line is approximately equal to the rapidity  $y$  for  $p \gg m$  and  $\theta \gg 1/\gamma$ , and in any case can be measured when the mass and momentum of the particle are unknown. From the definition one can obtain the identities

$$\sinh \eta = \cot \theta , \quad \cosh \eta = 1/\sin \theta , \quad \tanh \eta = \cos \theta . \quad (47.48)$$

**47.5.3. Partial waves :** The amplitude in the center of mass for elastic scattering of spinless particles may be expanded in Legendre polynomials

$$f(k, \theta) = \frac{1}{k} \sum_{\ell} (2\ell + 1) a_{\ell} P_{\ell}(\cos \theta) , \quad (47.49)$$

where  $k$  is the c.m. momentum,  $\theta$  is the c.m. scattering angle,  $a_{\ell} = (\eta_{\ell} e^{2i\delta_{\ell}} - 1)/2i$ ,  $0 \leq \eta_{\ell} \leq 1$ , and  $\delta_{\ell}$  is the phase shift of the  $\ell^{\text{th}}$  partial wave. For purely elastic scattering,  $\eta_{\ell} = 1$ . The differential cross section is

$$\frac{d\sigma}{d\Omega} = |f(k, \theta)|^2 . \quad (47.50)$$

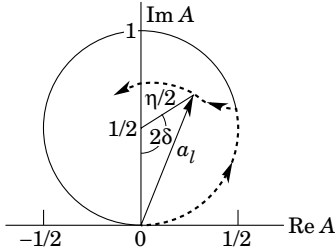
The optical theorem states that

$$\sigma_{\text{tot}} = \frac{4\pi}{k} \text{Im } f(k, 0) , \quad (47.51)$$

and the cross section in the  $\ell^{\text{th}}$  partial wave is therefore bounded:

$$\sigma_{\ell} = \frac{4\pi}{k^2} (2\ell + 1) |a_{\ell}|^2 \leq \frac{4\pi(2\ell + 1)}{k^2} . \quad (47.52)$$

The evolution with energy of a partial-wave amplitude  $a_{\ell}$  can be displayed as a trajectory in an Argand plot, as shown in Fig. 47.7.



**Figure 47.7:** Argand plot showing a partial-wave amplitude  $a_{\ell}$  as a function of energy. The amplitude leaves the unitary circle where inelasticity sets in ( $\eta_{\ell} < 1$ ).

The usual Lorentz-invariant matrix element  $\mathcal{M}$  (see Sec. 47.3 above) for the elastic process is related to  $f(k, \theta)$  by

$$\mathcal{M} = -8\pi\sqrt{s} f(k, \theta) , \quad (47.53)$$

so

$$\sigma_{\text{tot}} = -\frac{1}{2p_{\text{lab}} m_2} \text{Im } \mathcal{M}(t=0) , \quad (47.54)$$

where  $s$  and  $t$  are the center-of-mass energy squared and momentum transfer squared, respectively (see Sec. 47.4.1).

**47.5.3.1. Resonances:** The Breit-Wigner (nonrelativistic) form for an elastic amplitude  $a_{\ell}$  with a resonance at c.m. energy  $E_R$ , elastic width  $\Gamma_{\text{el}}$ , and total width  $\Gamma_{\text{tot}}$  is

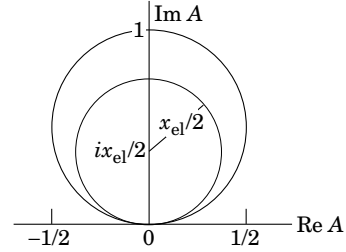
$$a_{\ell} = \frac{\Gamma_{\text{el}}/2}{E_R - E - i\Gamma_{\text{tot}}/2} , \quad (47.55)$$

where  $E$  is the c.m. energy. As shown in Fig. 47.8, in the absence of background the elastic amplitude traces a counterclockwise circle with center  $ix_{\text{el}}/2$  and radius  $x_{\text{el}}/2$ , where the elasticity  $x_{\text{el}} = \Gamma_{\text{el}}/\Gamma_{\text{tot}}$ . The amplitude has a pole at  $E = E_R - i\Gamma_{\text{tot}}/2$ .

The spin-averaged Breit-Wigner cross section for a spin- $J$  resonance produced in the collision of particles of spin  $S_1$  and  $S_2$  is

$$\sigma_{BW}(E) = \frac{(2J+1)}{(2S_1+1)(2S_2+1)} \frac{\pi}{k^2} \frac{B_{\text{in}} B_{\text{out}} \Gamma_{\text{tot}}^2}{(E - E_R)^2 + \Gamma_{\text{tot}}^2/4} , \quad (47.56)$$

where  $k$  is the c.m. momentum,  $E$  is the c.m. energy, and  $B_{\text{in}}$  and  $B_{\text{out}}$  are the branching fractions of the resonance into the entrance and exit channels. The  $2S+1$  factors are the multiplicities of the incident spin states, and are replaced by 2 for photons. This expression is valid only for an isolated state. If the width is not small,  $\Gamma_{\text{tot}}$  cannot be treated as a constant independent of  $E$ . There are many other forms for  $\sigma_{BW}$ , all of which are equivalent to the one given here in the narrow-width case. Some of these forms may be more appropriate if the resonance is broad.



**Figure 47.8:** Argand plot for a resonance.

The relativistic Breit-Wigner form corresponding to Eq. (47.55) is:

$$a_{\ell} = \frac{-m\Gamma_{\text{el}}}{s - m^2 + im\Gamma_{\text{tot}}} . \quad (47.57)$$

A better form incorporates the known kinematic dependences, replacing  $m\Gamma_{\text{tot}}$  by  $\sqrt{s}\Gamma_{\text{tot}}(s)$ , where  $\Gamma_{\text{tot}}(s)$  is the width the resonance particle would have if its mass were  $\sqrt{s}$ , and correspondingly  $m\Gamma_{\text{el}}$  by  $\sqrt{s}\Gamma_{\text{el}}(s)$  where  $\Gamma_{\text{el}}(s)$  is the partial width in the incident channel for a mass  $\sqrt{s}$ :

$$a_{\ell} = \frac{-\sqrt{s}\Gamma_{\text{el}}(s)}{s - m^2 + i\sqrt{s}\Gamma_{\text{tot}}(s)} . \quad (47.58)$$

For the  $Z$  boson, all the decays are to particles whose masses are small enough to be ignored, so on dimensional grounds  $\Gamma_{\text{tot}}(s) = \sqrt{s}\Gamma_0/m_Z$ , where  $\Gamma_0$  defines the width of the  $Z$ , and  $\Gamma_{\text{el}}(s)/\Gamma_{\text{tot}}(s)$  is constant. A full treatment of the line shape requires consideration of dynamics, not just kinematics. For the  $Z$  this is done by calculating the radiative corrections in the Standard Model.

## 47.6. Transverse variables

At hadron colliders, a significant and unknown proportion of the energy of the incoming hadrons in each event escapes down the beam-pipe. Consequently if invisible particles are created in the final state, their net momentum can only be constrained in the plane transverse to the beam direction. Defining the  $z$ -axis as the beam direction, this net momentum is equal to the missing transverse energy vector

$$\mathbf{E}_T^{\text{miss}} = -\sum_i \mathbf{p}_T(i) , \quad (47.59)$$

where the sum runs over the transverse momenta of all visible final state particles.

### 47.6.1. Single production with semi-invisible final state :

Consider a single heavy particle of mass  $M$  produced in association with visible particles which decays as in Fig. 47.1 to two particles, of which one (labeled particle 1) is invisible. The mass of the parent particle can be constrained with the quantity  $M_T$  defined by

$$\begin{aligned} M_T^2 &\equiv [E_T(1) + E_T(2)]^2 - [\mathbf{p}_T(1) + \mathbf{p}_T(2)]^2 \\ &= m_1^2 + m_2^2 + 2[E_T(1)E_T(2) - \mathbf{p}_T(1) \cdot \mathbf{p}_T(2)] , \end{aligned} \quad (47.60)$$

where

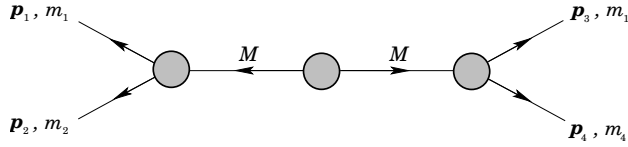
$$\mathbf{p}_T(1) = \mathbf{E}_T^{\text{miss}} . \quad (47.61)$$

This quantity is called the ‘transverse mass’ by hadron collider experimentalists but it should be noted that it is quite different from that used in the description of inclusive reactions [Eq. (47.39)]. The distribution of event  $M_T$  values possesses an end-point at  $M_T^{\text{max}} = M$ . If  $m_1 = m_2 = 0$  then

$$M_T^2 = 2|\mathbf{p}_T(1)||\mathbf{p}_T(2)|(1 - \cos \phi_{12}) , \quad (47.62)$$

where  $\phi_{ij}$  is defined as the angle between particles  $i$  and  $j$  in the transverse plane.

## 47.6.2. Pair production with semi-invisible final states :



**Figure 47.9:** Definitions of variables for pair production of semi-invisible final states. Particles 1 and 3 are invisible while particles 2 and 4 are visible.

Consider two identical heavy particles of mass  $M$  produced such that their combined center-of-mass is at rest in the transverse plane (Fig. 47.9). Each particle decays to a final state consisting of an invisible particle of fixed mass  $m_1$  together with an additional visible particle.  $M$  and  $m_1$  can be constrained with the variables  $M_{T2}$  and  $M_{CT}$  which are defined in Refs. 4 and 5.

**References:**

1. See, for example, J.J. Sakurai, *Modern Quantum Mechanics*, Addison-Wesley (1985), p. 172, or D.M. Brink and G.R. Satchler, *Angular Momentum*, 2nd ed., Oxford University Press (1968), p. 20.
2. I. Hinchliffe *et al.*, Phys. Rev. **D55**, 5520 (1997).
3. B.C. Allanach *et al.*, JHEP **0009**, 004 (2000).
4. C.G. Lester and D.J. Summers, Phys. Lett. **B463**, 99 (1999).
5. D.R. Tovey, JHEP **0804**, 034 (2008).

## 48. RESONANCES

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### 48.1. General Considerations

Perturbative methods can be applied to systems of quarks and gluons only for large momentum transfers (see review on 'Quantum chromodynamics') and, under certain conditions, to some properties of systems that contain heavy quarks (see review on 'Heavy-Quark and Soft-Collinear Effective Theory'). In general, however, dealing with QCD in the low momentum transfer region is a very complicated, non-perturbative problem where quarks and gluons are confined within color neutral hadrons. Physical states show up as poles of the  $S$ -matrix either on the physical sheet (bound states) or on the unphysical sheets (resonances) and manifest themselves as structures in experimental observables.

Resonances can show up either in so-called formation experiments typically of the kind

$$A + B \rightarrow R \rightarrow C_1 + \dots + C_n ,$$

where they become visible in an energy scan (a perfect example of this being the  $R$ -function measured in  $e^+e^-$  annihilations — *cf.* the corresponding plots in the review on 'Plots of Cross sections and related quantities'), or together with a spectator particle  $S$  in production experiments of the kind

$$A \rightarrow R + S \rightarrow [C_1 + \dots + C_n] + S .$$

In the latter case the resonances properties are commonly extracted from a Dalitz plot analysis (see review on 'Kinematics') or projections thereof. Multi-particle final states are often parametrized in terms of successive decays of two-body resonances.

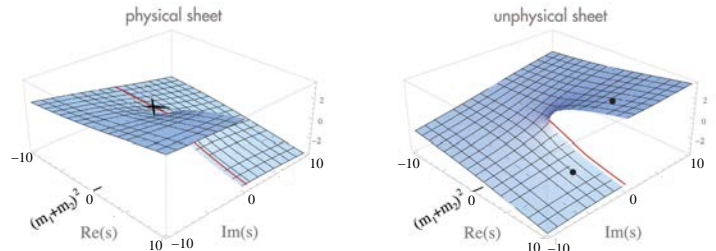
Resonance phenomena are very rich: while typical hadronic widths are of the order of 100 MeV (e.g., for the meson resonances  $\rho(770)$  or  $\psi(4040)$  or the baryon resonance  $\Delta(1232)$ ) corresponding to a life time of  $10^{-23}$  s, the widths can also be as small as sub MeV (e.g. of  $X(3872)$ ) or as large as several hundred MeV (e.g. of the meson resonances  $f_0(500)$  or  $D_1(2430)$  or the baryon resonance  $N(2190)$ ).

Ideally a resonance appears as a peak in the total cross section. If the structure is narrow and if there are no relevant thresholds or other resonances nearby, the resonance properties may be extracted employing a standard Breit-Wigner parametrization if necessary improved by using an energy dependent width term (*cf.* Sec. 2.1 of this review). However, in general unitarity and analyticity call for the use of more refined tools, e.g. when there are overlapping resonances with the same quantum numbers the resonance terms should not simply be added but combined in a non-trivial way either in a  $K$ -matrix approximation (*cf.* Sec. 2.3 of this review) or using more refined methods (*cf.* Sec. 1.4 of this review). Only then the proper pole parameters can be extracted that are universal resonance properties — on the contrary, Breit-Wigner parameters are typically reaction dependent. In addition, for broad resonances there is no direct relation anymore between pole location and the total width/life time — then the pole residues need to be used in order to quantify the decay properties of a given state (*cf.* Sec. 3 of this review).

For simplicity, throughout this review the formulas are given for distinguishable, scalar particles. The additional complications that appear in the presence of spins can be controlled in the helicity framework developed by Jacob and Wick [1], or in a non-relativistic [2] or relativistic [3] tensor operator formalism. Within these frames, sequential (cascade) decays are commonly treated as a coherent sum of two-body interactions. Therefore below most explicit expressions are given for two-body kinematics.

#### 48.1.1. Properties of the $S$ -matrix :

The unitary operator that connects asymptotic *in* and *out* states is called the  $S$ -matrix. It is an analytic function in the Mandelstam plane up to its branch points and poles. Branch points appear whenever there is a channel opening — at each threshold for massive particles the number of Riemann sheets doubles. Poles refer either to bound states or to resonances. The former poles are located on the physical sheet, the latter are located on the unphysical sheet closest to the physical one, often called the second sheet; each can be accompanied by mirror poles. If there are resonances in subsystems of multi-particle states, branch points appear in the complex plane of the unphysical sheet(s). Any of these singularities leads to some structure in the observables (see also Ref. [4]). In a partial wave decomposed amplitude additional singularities not related to resonance physics may emerge as a result of the partial-wave projection. For a discussion see, e.g., Ref. [5].



**Figure 48.1:** Sketch of the imaginary part of a typical single-channel amplitude in the complex  $s$ -plane. The solid dots indicate allowed positions for resonance poles, the cross for a bound state. The solid line is the physical axis (shifted by  $i\epsilon$  into the physical sheet). The two sheets are connected smoothly along their discontinuities.

If for simplicity we now restrict ourselves to reactions involving four particles, the kinematics of the reaction are fully described by the Mandelstam variables  $s$ ,  $t$  and  $u$ , only two of them being independent (*cf.* Eqs. (28)-(31) of the kinematics review). Bound state poles are allowed only on the real  $s$ -axis below the lowest threshold. There is no restriction for the location of poles on the unphysical sheets — only that analyticity requires that, if there is a pole at some complex value of  $s$ , there must be another pole at its complex conjugate value,  $s^*$ . The pole with a negative imaginary part is closer to the physical axis and thus influences the observables in the vicinity of the resonance region more strongly, however, at the threshold both poles are always equally important. This is illustrated in Fig. 48.1.

The  $S$ -matrix is related to the scattering matrix  $\mathcal{M}$  (*c.f.* Eq. (8) of the kinematics review). For two-body scattering it can be cast into the form

$$S_{ab} = I_{ab} - 2i\sqrt{\rho_a}\mathcal{M}_{ab}\sqrt{\rho_b} . \quad (48.1)$$

$\mathcal{M}$  is a matrix in channel space and depends, for two-body scattering, on both  $s$  and  $t$ . The channel indices  $a$  and  $b$  are multi-indices specifying all properties of the channel including the conserved quantum numbers. The two-body phase-space  $\rho$  is given (*cf.* Eq. 12 of the kinematics review) by

$$\rho_a(s) = \frac{1}{16\pi} \frac{2|\vec{q}_a|}{\sqrt{s}} . \quad (48.2)$$

with  $q_a$  denoting the relative momentum of the decay particles of channel  $a$ , with masses  $m_1$  and  $m_2$ , *cf.* Eq. (20a) of the kinematics review.

As discussed below, unitarity puts strong constraints on the scattering matrix. Further constraints come, e.g., from crossing symmetry and duality [6].

### 48.1.2. Consequences from unitarity :

In what follows, scattering amplitudes  $\mathcal{M}$  and decay amplitudes  $\mathcal{A}$  will be distinguished, since unitarity puts different constraints on these. The discontinuity of the scattering amplitude from channel  $a$  to channel  $b$  [7] is constrained by unitarity to

$$i[\mathcal{M}_{ba} - \mathcal{M}_{ab}^*] = (2\pi)^4 \sum_c \int d\Phi_c \mathcal{M}_{cb}^* \mathcal{M}_{ca} . \quad (48.3)$$

Using  $\text{Disc}(\mathcal{M}(s)) = 2i \text{Im}(\mathcal{M}(s + i\epsilon))$  the optical theorem follows

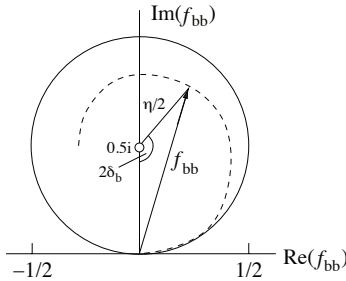
$$\text{Im}(\mathcal{M}_{aa}|_{\text{forward}}) = 2q_a \sqrt{s} \sigma_{\text{tot}}(a \rightarrow \text{anything}) . \quad (48.4)$$

The unitarity relation for a decay amplitude of a heavy state  $H$  into a channel  $a$  is given by

$$i[\mathcal{A}_a^H - \mathcal{A}_a^{H*}] = (2\pi)^4 \sum_c \int d\Phi_c \mathcal{M}_{ca}^* \mathcal{A}_c^H . \quad (48.5)$$

From Eq. (48.5) Watson's theorem follows straightforwardly: the phase of  $\mathcal{A}$  agrees with that of  $\mathcal{M}$  as long as only a single channel contributes. For systems where the phase shifts are known like  $\pi\pi$  in  $S$ - and  $P$ -waves for low energies,  $\mathcal{A}^H$  can be calculated in a model-independent way using dispersion theory [8]. Those methods can also be generalized to three-body final states [9] and were applied to  $\eta \rightarrow \pi\pi\pi$  in Refs. [10,11,12] and to  $\phi$  and  $\omega$  to  $3\pi$  in Ref. [13].

### 48.1.3. Partial-wave decomposition :



**Figure 48.2:** Argand plot showing a diagonal element of a partial-wave amplitude,  $a_{bb}$ , as a function of energy. The amplitude leaves the unitary circle (solid line) as soon as inelasticity sets in,  $\eta < 1$  (dashed line).

In general, a physical amplitude  $\mathcal{M}$  (c.f. Eq. (8) of the kinematics review) is a matrix in channel space. It depends, for two-body scattering, on both  $s$  and  $t$ . It is often convenient to expand the amplitudes in partial waves. For this purpose one defines for the transition matrix from channel  $a$  to channel  $b$

$$\mathcal{M}_{ba}(s, t) = \sum_{L=0}^{\infty} (2L+1) \mathcal{M}_{ba}^L(s) P_L(\cos(\theta)) , \quad (48.6)$$

where  $L$  denotes the angular momentum—in the presence of spins the initial and final value of  $L$  does not need to be equal. To simplify notations below we will drop the label  $L$ . The function  $\mathcal{M}_{ba}(s)$  is expressed in terms of the partial-wave amplitudes  $f_{ba}(s)$  via

$$\mathcal{M}_{ba}(s) = -f_{ba}(s) / \sqrt{\rho_a \rho_b} . \quad (48.7)$$

The partial-wave amplitudes  $f_{ba}$  depend on  $s$  only. Using  $S_{ba} = \delta_{ba} + 2if_{ba}$  one gets from the unitarity of the  $S$ -matrix

$$f_{bb} = (\eta \exp(2i\delta_b) - 1) / 2i , \quad (48.8)$$

where  $\delta_b(\eta)$  denotes the phase shift (elasticity parameter — also called inelasticity) for the scattering from channel  $b$  to channel  $b$ . One has  $0 \leq \eta \leq 1$ , where  $\eta = 1$  refers to purely elastic scattering. The evolution with energy of a partial-wave amplitude  $f_{bb}$  can be displayed as a trajectory in an Argand plot, as shown in Fig. 48.2. In case of a two-channel problem the off-diagonal element is typically parametrized as  $f_{ba} = \sqrt{1 - \eta^2/2} \exp(i(\delta_b + \delta_a))$ .

### 48.1.4. Explicit parametrizations for scattering and production amplitudes :

It is often convenient to decompose the physical amplitude  $\mathcal{M}$  into a pole part and a non-pole part, often called background

$$\mathcal{M} = \mathcal{M}^{\text{b.g.}} + \mathcal{M}^{\text{pole}} . \quad (48.9)$$

The splitting given in Eq. (48.9) is reaction dependent and not unique (see, e.g., the discussion in Ref. [14]), such that some resonances show up differently in different reactions. What are independent of the reaction, however, are the location of the pole of a given resonance  $R$  in the complex  $s$ -plane,  $s_R$ , and its residues, or, more accurately, the pole couplings introduced in the last section of this review. Those parameters capture all the properties of a given resonance. The decomposition of Eq. (48.9) is employed, e.g., in Ref. [15] to study the lineshape of  $\psi(3770)$  and in Refs. [16,17] to investigate  $\pi N$  scattering. Traditionally one introduces the notation

$$\sqrt{s_R} = M_R - i\Gamma_R/2 , \quad (48.10)$$

where  $M_R$  and  $\Gamma_R$  are referred to as mass and total width of the resonance  $R$ , respectively. Note, the standard Breit-Wigner parameters  $M_{\text{BW}}$  and  $\Gamma_{\text{BW}}$ , also introduced below, in general deviate from the pole parameters, e.g., due to finite width effects and the influence of thresholds.

If there are  $N$  resonances in a particular channel,

$$\mathcal{M}_{ba}^{\text{pole}}(s) = \gamma_b(s) [1 - V^R(s) \Sigma(s)]_{bc}^{-1} V_{ca}^R(s) \gamma_a(s) . \quad (48.11)$$

where all ingredients are matrices in channel space. Especially

$$V_{ab}^R(s) = - \sum_{n=1}^N \frac{g_{nb} g_{na}}{s - M_n^2} , \quad (48.12)$$

$\gamma_a$  and  $\Sigma_a$  denote the normalized vertex function and the self-energy, respectively, while  $g_{na}$  denotes the coupling of the resonance  $R_n$  to channel  $a$  and  $M_n$  its mass parameter (not to be confused with the pole position). The sign in Eq. (48.12) is necessary to render the  $g$ -parameters real. A relation analogous to Eq. (48.5) holds for any kind of production amplitude — especially for the normalized vertex functions, however, with the final state interaction provided by  $\mathcal{M}^{\text{b.g.}}$

$$i[\gamma_a - \gamma_a^*] = (2\pi)^4 \sum_c d\Phi_c \left( \mathcal{M}^{\text{b.g.}} \right)_{ca}^* \gamma_c . \quad (48.13)$$

The discontinuity of the self-energy  $\Sigma_a(s)$  is

$$i[\Sigma_a - \Sigma_a^*] = (2\pi)^4 \int d\Phi_a |\gamma_a|^2 . \quad (48.14)$$

The real part of  $\Sigma_a$  can be calculated from Eq. (48.14) via a properly subtracted dispersion integral. If  $\mathcal{M}^{\text{b.g.}}$  is unitary, the use of Eq. (48.11) leads to a unitary full amplitude, cf. Eq. (48.9).

For a single resonance ( $N = 1$ ) Eq. (48.11) reads

$$\mathcal{M}_{ba}^{\text{pole}}(s)|_{N=1} = -\gamma_b(s) \frac{g_b g_a}{s - \hat{M}_R(s)^2 + i\sqrt{s}\Gamma_R(s)_{\text{tot}}} \gamma_a(s) , \quad (48.15)$$

where the mass function  $\hat{M}_R(s)^2 = M^2 + \sum_c g_c^2 \text{Re}(\Sigma_c)$ . The imaginary part of the self-energy gives the width of the resonance via

$$\Gamma_c^R(s) = \frac{(2\pi)^4}{2\sqrt{s}} g_c^2 \int d\Phi_c |\gamma_c|^2 ; \quad \Gamma^R(s)_{\text{tot}} = \sum_c \Gamma_c^R(s) . \quad (48.16)$$

Here the sum runs over all channels. Eq. (48.16) agrees with Eq. (10) of the kinematics review.

In the absence of left-hand cuts in the production mechanism, the decay amplitude  $\mathcal{A}^H$  can be written as

$$\mathcal{A}_a^H(s) = \gamma_a(s) \left[ 1 - V^R(s) \Sigma(s) \right]_{ab}^{-1} \mathcal{P}_b^H(s) , \quad (48.17)$$

where  $\mathcal{P}^H$  is a vector in channel space that may be parametrized as

$$\mathcal{P}_b^H(s) = p_b(s) - \sum_{n=1}^N \frac{g_{nb} \alpha_n^H}{s - M_n^2} \quad (48.18)$$

and the masses  $M_n$  need to agree with those in  $V_R$ . The function  $p_a(s)$  is a background term and the  $\alpha_n^H$  denote the coupling of the heavy state  $H$  to the particular resonance  $R_n$ . If there are additional particles in the final state of the studied decay of heavy state  $H$ , not included in the non-perturbative treatment of Eq. (48.17), then they also contain the corresponding kinematic factors related to their coupling. If these additional particles are interacting strongly, a complete few-body treatment of the final state becomes necessary, especially since rescattering effects can introduce additional complex phases [18]. However, in practice those effects as well as those from missing channels are often parametrized by choosing the parameters  $\alpha_n^H$  complex valued. With some additional assumptions, Eq. (48.9) and Eq. (48.17) were employed in Ref. [19] to study the pion vector form factor. An alternative parametrization for the production amplitude that is convenient, if the full matrix  $\mathcal{M}$  — including the resonances — is known, cf. Ref. [20]

$$\mathcal{A}_a^H(s) = \mathcal{M}_{ab}(s) \tilde{\mathcal{P}}_b^H(s). \quad (48.19)$$

The function  $\tilde{\mathcal{P}}_b^H(s)$  needs to cancel the left-hand cuts of  $\mathcal{M}$  and therefore could be strongly energy dependent. In actual applications a low-order polynomial turned out to be sufficient — c.f. Ref. [21,22] for a study of  $\gamma\gamma \rightarrow \pi\pi$ . As above, to preserve unitarity the coefficients of  $\tilde{\mathcal{P}}_b^H(s)$  need to be real, however, in practice rescattering effects or missing channels are parametrized by complex valued parameters.

Three-body decays are often represented by Dalitz plots. It is often of interest to quantify the contribution of a single amplitude  $\mathcal{A}_a^H$  to the decay of a heavy resonance  $H$ , where now  $\mathcal{A}_a^H$  needs to be generalized to three body kinematics either completely by considering the full three-body final state interactions or effectively by choosing complex vertex parameters. Then fractional contributions are introduced (since different intermediate states leading to the same final state interfere, the assignment of branching ratios is to be taken with some caution) via

$$F_a^H = \frac{\int d\Phi |\mathcal{A}_a^H|^2}{\int d\Phi |\sum_a \mathcal{A}_a^H|^2} \quad (48.20)$$

where the phase space integral  $d\Phi$  extends over the Dalitz plot region and the angular dependence of the subsystems needs to be kept (cf. Eq. (48.6)). Typically the effect of interference terms in the denominator is small.

The formulas given so far are completely general. However, they require as input, e.g., information on the non-resonant scattering in the various channels. It is therefore often necessary and appropriate to find approximations/parametrizations.

## 48.2. Common parametrizations for resonances

In most common parametrizations the non-pole interaction,  $\mathcal{M}^{\text{b.g.}}$ , is omitted. While this is a bad approximation for, e.g., scalar-isoscalar  $\pi\pi$  interactions at very low energies [23], under more favorable conditions this can be justified. Thus in what follows we will assume  $\mathcal{M}^{\text{b.g.}} = 0$ , which leads to real vertex functions. For two-body channels one writes

$$\gamma(s)_a = q_a^{L_a} F_{L_a}(q_a, q_0),$$

where  $L_a$  denotes the angular momentum of the decay products, giving rise to the centrifugal barrier  $q_a^{L_a}$ , where  $q_a$  denotes the relative momentum of the outgoing particle pair defined in the rest frame of the decaying particle, cf. Eq. (20a) of the kinematics review. Often one introduces a phenomenological form factor, here denoted by  $F_{L_a}(q_a, q_0)$ . It depends on the channel momentum as well as some intrinsic scale  $q_0$ . Often the Blatt-Weisskopf form is chosen [24,25], where, e.g.,  $F_0^2 = 1$ ,  $F_1^2 = 2/(q_a + q_0)$  and  $F_2^2 = 13/((q_a - 3q_0)^2 + 9q_a q_0)$ . In addition, for isolated, narrow resonances the couplings  $g_a$  can be related to the partial widths,  $\Gamma_{R \rightarrow a}$ , via

$$g_a = \frac{1}{\gamma_a(s_R)} \sqrt{\frac{M_R \Gamma_{R \rightarrow a}}{\rho_a}}, \quad (48.21)$$

where  $M_R$  was defined in Eq. (48.10).

### 48.2.1. The Breit-Wigner and Flatté Parametrizations :

If there is only a single resonance present and all relevant thresholds are far away, then one may replace  $\Gamma_R(s)_{\text{tot}}$  with a constant,  $\Gamma_{\text{BW}}$ . Under these conditions also the real part of  $\Sigma$  is a constant that can be absorbed into the mass parameter and Eq. (48.15) simplifies to

$$\mathcal{M}_{ba}^{\text{pole}} \Big|_{N=1} = - \frac{g_b g_a}{s - M_{\text{BW}}^2 + i\sqrt{s}\Gamma_{\text{BW}}}, \quad (48.22)$$

which is the standard Breit-Wigner parametrization. For a narrow resonance it is common to replace  $\sqrt{s}$  by  $M_{\text{BW}}$ . If there are nearby relevant thresholds,  $\Gamma_{\text{BW}}$  needs to be replaced by  $\Gamma(s)$ . For two-body decays one writes

$$\Gamma(s) = \sum_c \Gamma_{R \rightarrow c} \left( \frac{q_c}{q_{Rc}} \right)^{2L_c+1} \left( \frac{M_R}{\sqrt{s}} \right) \left( \frac{F_{L_c}(q_c, q_0)}{F_{L_c}(q_{Rc}, q_0)} \right)^2, \quad (48.23)$$

where  $q_{Rc} = q(M_{\text{BW}})_c$  denotes the decay momentum of resonance  $R$  into channel  $c$ . The Breit-Wigner parameters  $M_{\text{BW}}$  and  $\Gamma_{\text{BW}}$  agree with the pole parameters only if  $M_R \Gamma(M_R) \ll M_{\text{thr.}}^2 - M_R^2$ , with  $M_{\text{thr.}}$  for the closest relevant threshold. Otherwise the Breit-Wigner parameters deviate from the pole parameters and are reaction dependent.

If there is more than one resonance in one partial wave that significantly couples to the same channels, it is in general incorrect to use a sum of Breit-Wigner functions, for it may violate unitarity constraints. Then more refined methods should be used, like the  $K$ -matrix approximation described in the next section.

Below the corresponding threshold,  $q_c$  in Eq. (48.23) must be continued analytically: if, e.g., the particles in channel  $c$  have equal mass  $m_c$ , then

$$q_c = \frac{i}{2} \sqrt{4m_c^2 - s} \quad \text{for } \sqrt{s} < 2m_c. \quad (48.24)$$

The resulting line shape above and below the threshold of channel  $c$  is called Flatté parametrization [26]. If the coupling of a resonance to the channel opening nearby is very strong, the Flatté parametrization shows a scaling invariance and does not allow for an extraction of individual partial decay widths, but only of ratios [27].

### 48.2.2. The $K$ -matrix approximation :

As soon as there is more than one resonance in one channel, the use of the  $K$ -matrix approximation should be preferred compared to the Breit-Wigner parametrization discussed above. From the considerations formulated in Eq. (48.11), the  $K$ -matrix approximation follows straightforwardly by replacing the self-energy  $\Sigma_c$  by its imaginary part in the absence of  $\mathcal{M}^{\text{b.g.}}$ , but keeping the full matrix structure of  $V^R$ . Thus, for two-body intermediate states one writes within this scheme for the self-energy

$$\Sigma(s)_c \rightarrow i\rho_c \gamma(s)_c^2. \quad (48.25)$$

However, in distinction to the Breit-Wigner approach,  $V^R$ , then called  $K$ -matrix, is kept in the form of Eq. (48.12). The decay amplitude given in Eq. (48.17) then takes the form of the standard  $P$ -vector formalism introduced in Ref. [28]. For  $N = 1$  the amplitude derived from the  $K$ -matrix is identical to that of Eq. (48.22).

Some authors use the analytic continuation of  $\rho_c$  below the threshold via the analytic continuation of the particle momentum as described above [29,30].

### 48.2.3. Further improvements :

The  $K$ -matrix described above usually allows one to get a proper fit of physical amplitudes and it is easy to deal with, however, it also has an important deficit: it violates constraints from analyticity — e.g.,  $\rho_a$ , defined in Eq. (48.2), is ill-defined at  $s = 0$  and for unequal masses develops an unphysical cut. In addition, the analytic continuation of the amplitudes into the complex plane is not controlled and typically the parameters of broad resonances come out wrong (see, e.g., minireview on scalar mesons). A method to improve the analytic properties was suggested in Refs. [31,32,33,34]. It basically

amounts to replacing the phase-space factor  $i\rho_a$  in Eq. (48.25) by an analytic function that produces the identical imaginary part on the right-hand cut. In the simplest case of a channel with equal masses the expressions that can be used for real values of  $s$  read

$$-\frac{\hat{\rho}_a}{\pi} \log \left| \frac{1+\hat{\rho}_a}{1-\hat{\rho}_a} \right|, -\frac{2\hat{\rho}_a}{\pi} \arctan \left( \frac{1}{\hat{\rho}_a} \right), -\frac{\hat{\rho}_a}{\pi} \log \left| \frac{1+\hat{\rho}_a}{1-\hat{\rho}_a} \right| + i\hat{\rho}_a$$

for  $s < 0$ ,  $0 < s < 4m_a^2$ , and  $4m_a^2 < s$ , respectively, with  $\hat{\rho}_a = \sqrt{1-4m_a^2/s}$  for all values of  $s$ , extending the expression of Eq. (48.2) into the regime below threshold. The more complicated expression for the case of different masses can be found, e.g., in Ref. [32].

If there is only a single resonance in a given channel, it is possible to feed the imaginary part of the Breit-Wigner function, Eq. (48.22) with an energy-dependent width, directly into a dispersion integral to get a resonance propagator with the correct analytic structure [35,36].

### 48.3. Properties of resonances

A resonance is characterized not only by its complex pole position but also by its residues that quantify its couplings to the various channels and allow one to define a branching ratio also for broader resonances. In the Meson Particle Listings the two-photon width of  $f_0(500)$  is defined in terms of the corresponding residue. The Baryon Particle Listings give the elastic pole residues and normalized transition residues. However, different conventions are used in the two sectors, which are shortly outlined here.

In the close vicinity of a pole the scattering matrix  $\mathcal{M}$  can be written as

$$\lim_{s \rightarrow s_R} \mathcal{M}_{ba} = -\frac{\mathcal{R}_{ba}}{s - s_R}, \quad (48.26)$$

where  $s_R$  denotes the pole position of the resonance R. The sign convention in Eq. (48.26) is consistent with that of Eq. (48.12). The residues may be calculated via an integration along a closed contour around the pole using

$$\mathcal{R}_{ba} = \frac{i}{2\pi} \oint ds \mathcal{M}_{ba}.$$

The factorization of the residue  $(\mathcal{R}_{ba})^2 = \mathcal{R}_{aa} \times \mathcal{R}_{bb}$  allows one to introduce pole couplings according to

$$\tilde{g}_a = \mathcal{R}_{ba} / \sqrt{\mathcal{R}_{bb}}. \quad (48.27)$$

The pole couplings are the only quantities that allow one to quantify the transition strength of a given resonance to some channel  $a$  independent of how the particular resonance was produced. For a single, narrow state with an energy-independent background in the resonance region, far away from all relevant thresholds one finds  $\tilde{g}_a = \gamma_a(s_R)g_a$  with the real valued resonance couplings  $g_a$  defined in Eq. (48.12) accompanied by the complex valued vertex functions  $\gamma_a$  introduced in Eq. (48.11). Based on this observation one may use the straightforward generalization of Eq. (48.21) to define a partial width and a branching fraction even for a broad resonance via

$$\Gamma_{R \rightarrow a} = \frac{|\tilde{g}_a|^2}{M_R} \rho_a(M_R^2) \quad \text{and} \quad Br_a = \Gamma_{R \rightarrow a} / \Gamma_R, \quad (48.28)$$

where  $M_R$  and  $\Gamma_R$  were introduced in Eq. (48.10). This expression was used to define a two-photon width for the broad  $f_0(500)$  (also called  $\sigma$ ) [21,22]. Eq. (48.28) defines a partial decay width independent of the reaction used to extract the parameters. It maps smoothly onto the standard definitions for narrow resonances — cf. Eq. (48.16). There are cases where a resonance couples to a channel that opens only above  $M_R$ . A prominent example for this being  $f_0(980)$  to  $\bar{K}K$ . If one wants to define a branching fraction that also captures this situation one may define

$$Br'_a = \int_{\text{threshold}}^{\infty} \frac{ds}{\pi} \frac{|\tilde{g}_a|^2 \rho(s)}{|D(s)|^2}. \quad (48.29)$$

Here one needs to assume a line shape for the resonance R. A possible choice is a Flatté form  $|D(s)|^2 = (M_R^2 - s)^2 + (\sum_a |\tilde{g}_a|^2 \rho_a(s))^2$ . The only model-independent quantities are the pole couplings/residues — both forms, Eq. (48.28) and Eq. (48.29), are in general not directly related to observables but meant to quantify the effect of the pole couplings by employing better known quantities.

In the baryon sector it is common to define the residue with respect to the partial-wave amplitudes  $f_{ba}(s)$  defined in Eq. (48.7) and with respect to  $\sqrt{s}$  instead of  $s$ . The two definitions are related via

$$Res(a \rightarrow b) = -\sqrt{\frac{\rho_a(s_R)\rho_b(s_R)}{4s_R}} \mathcal{R}_{ba}, \quad (48.30)$$

where the phase space factors are to be evaluated at the pole. The elastic pole residues for  $a \rightarrow a$  scattering, in the baryon listings called  $r$ , are

$$r = -Res(a \rightarrow a). \quad (48.31)$$

One may now define the partial decay widths and the branching ratios of a resonance R into channel  $a$  at its pole position on the basis of the residues introduced in Eq. (48.30)

$$\Gamma_{R \rightarrow a} = 2 |Res(a \rightarrow b)| \quad \text{and} \quad BR_a = 2 |Res(a \rightarrow a)| / \Gamma_R. \quad (48.32)$$

The only difference between the definitions of the branching ratio of Eq. (48.28) and Eq. (48.32) is that for the former the phase space factors are evaluated on the real axis while for the latter they are evaluated at the pole. The Baryon Particle Listings give information on the  $\pi N$  elastic residues,  $r$ , on various normalized  $\pi N \rightarrow a$  transition residues, and on branching ratios.

#### References:

1. M. Jacob and G.C. Wick, Ann. Phys. **7**, 404 (1959) [Ann. Phys. **281**, 774 (2000)].
2. C. Zemach, Phys. Rev. **140B**, 97 (1965); Phys. Rev. **140B**, 109 (1965).
3. A.V. Anisovich *et al.*, J. Phys. **G28**, 15 (2002); Eur. Phys. J. **A24**, 111 (2005).
4. A rapid change in an amplitude is not an unambiguous signal of a singularity of the  $S$ -matrix [37], however, for realistic interactions this connection holds.
5. G. Höhler, *Pion-Nucleon Scattering – Methods and Results of Phenomenological Analyses*, Springer-Verlag Berlin, Heidelberg, New York, 1983.
6. M. Fukugita and K. Igi, Phys. Reports **31**, 237 (1977).
7. M.P. Peskin and D.V. Schroeder, *An Introduction to Quantum Field Theory*, Westview Press, 1995.
8. R. Omnès, Nuovo Cimento **8**, 316 (1958).
9. N.N. Khuri and S.B. Treiman, Phys. Rev. **119**, 1115 (1960).
10. J. Kambor, C. Wiesendanger, and D. Wyler, Nucl. Phys. **B465**, 215 (1996).
11. A.V. Anisovich and H. Leutwyler, Phys. Lett. **B375**, 335 (1996).
12. S.P. Schneider, B. Kubis, and C. Ditsche, JHEP **1102**, 028 (2011).
13. F. Niecknig, B. Kubis, and S.P. Schneider, Eur. Phys. J. **C72**, 2014 (2012).
14. D. Djukanovic, J. Gegelia, and S. Scherer, Phys. Rev. **D76**, 037501 (2007).
15. N.N. Achasov and G.N. Shestakov, Phys. Rev. **D86**, 114013 (2012).
16. A. Matsuyama, T. Sato, and T.-S.H. Lee, Phys. Reports **439**, 193 (2007).
17. M. Döring *et al.*, Phys. Lett. **B681**, 26 (2009).
18. I. Caprini, Phys. Lett. **B638**, 468 (2006).
19. C. Hanhart, Phys. Lett. **B715**, 170 (2012).
20. K.L. Au, D. Morgan, and M. R. Pennington, Phys. Rev. **D35**, 1633 (1987).
21. D. Morgan and M.R. Pennington, Z. Phys. **C37**, 431 (1988) [Erratum: Z. Phys. **C39**, 590 (1988)].
22. D. Morgan and M.R. Pennington, Z. Phys. **C48**, 623 (1990).
23. J. Gasser and U.G. Meißner, Nucl. Phys. **B357**, 90 (1991).



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24. J. Blatt and V. Weisskopf, *Theoretical Nuclear Physics*, New York: John Wiley & Sons (1952).
  25. S.U. Chung *et al.*, Ann. Phys. **4**, 404 (1995).
  26. S.M. Flatté, Phys. Lett. **B63**, 224 (1976).
  27. V. Baru *et al.*, Eur. Phys. J. **A23**, 523 (2005).
  28. I.J.R. Aitchison, Nucl. Phys. **A189**, 417 (1972).
  29. J.H. Reid and N.N. Trofimenkoff, J. Math. Phys. **25**, 3540 (1984).
  30. V.V. Anisovich and A.V. Sarantsev, Eur. Phys. J. **A16**, 229 (2003).
  31. M.R. Pennington *et al.*, Eur. Phys. J. **C56**, 1 (2008).
  32. J.A. Oller and E. Oset, Phys. Rev. **D60**, 074023 (1999).
  33. N.N. Achasov and A.V. Kiselev, Phys. Rev. **D83**, 054008 (2011).
  34. A.V. Anisovich *et al.*, Phys. Rev. **D84**, 076001 (2011).
  35. E. L. Lomon and S. Pacetti, Phys. Rev. **D85**, 113004 (2012) [Erratum: Phys. Rev. **D86**, 039901 (2012)].
  36. B. Moussallam, Eur. Phys. J. **C73**, 2539 (2013).
  37. G. Calucci, L. Fonda, and G.C. Ghirardi, Phys. Rev. **166**, 1719 (1968).

## 49. CROSS-SECTION FORMULAE FOR SPECIFIC PROCESSES

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### PART I: STANDARD MODEL PROCESSES

Setting aside leptonproduction (for which, see Sec. 16 of this *Review*), the cross sections of primary interest are those with light incident particles,  $e^+e^-$ ,  $\gamma\gamma$ ,  $q\bar{q}$ ,  $gq$ ,  $gg$ , etc., where  $g$  and  $q$  represent gluons and light quarks. The produced particles include both light particles and heavy ones -  $t$ ,  $W$ ,  $Z$ , and the Higgs boson  $H$ . We provide the production cross sections calculated within the Standard Model for several such processes.

#### 49.1. Resonance Formation

Resonant cross sections are generally described by the Breit-Wigner formula (Sec. 19 of this *Review*).

$$\sigma(E) = \frac{2J+1}{(2S_1+1)(2S_2+1)} \frac{4\pi}{k^2} \left[ \frac{\Gamma^2/4}{(E-E_0)^2 + \Gamma^2/4} \right] B_{in} B_{out}, \quad (49.1)$$

where  $E$  is the c.m. energy,  $J$  is the spin of the resonance, and the number of polarization states of the two incident particles are  $2S_1+1$  and  $2S_2+1$ . The c.m. momentum in the initial state is  $k$ ,  $E_0$  is the c.m. energy at the resonance, and  $\Gamma$  is the full width at half maximum height of the resonance. The branching fraction for the resonance into the initial-state channel is  $B_{in}$  and into the final-state channel is  $B_{out}$ . For a narrow resonance, the factor in square brackets may be replaced by  $\pi\Gamma\delta(E-E_0)/2$ .

#### 49.2. Production of light particles

The production of point-like, spin-1/2 fermions in  $e^+e^-$  annihilation through a virtual photon,  $e^+e^- \rightarrow \gamma^* \rightarrow f\bar{f}$ , at c.m. energy squared  $s$  is given by

$$\frac{d\sigma}{d\Omega} = N_c \frac{\alpha^2}{4s} \beta [1 + \cos^2\theta + (1-\beta^2)\sin^2\theta] Q_f^2, \quad (49.2)$$

where  $\beta$  is  $v/c$  for the produced fermions in the c.m.,  $\theta$  is the c.m. scattering angle, and  $Q_f$  is the charge of the fermion. The factor  $N_c$  is 1 for charged leptons and 3 for quarks. In the ultrarelativistic limit,  $\beta \rightarrow 1$ ,

$$\sigma = N_c Q_f^2 \frac{4\pi\alpha^2}{3s} = N_c Q_f^2 \frac{86.8 \text{ nb}}{s (\text{GeV}^2)}. \quad (49.3)$$

The cross section for the annihilation of a  $q\bar{q}$  pair into a distinct pair  $q'\bar{q}'$  through a gluon is completely analogous up to color factors, with the replacement  $\alpha \rightarrow \alpha_s$ . Treating all quarks as massless, averaging over the colors of the initial quarks and defining  $t = -s \sin^2(\theta/2)$ ,  $u = -s \cos^2(\theta/2)$ , one finds [1]

$$\frac{d\sigma}{d\Omega}(q\bar{q} \rightarrow q'\bar{q}') = \frac{\alpha_s^2}{9s} \frac{t^2 + u^2}{s^2}. \quad (49.4)$$

Crossing symmetry gives

$$\frac{d\sigma}{d\Omega}(qq' \rightarrow qq') = \frac{\alpha_s^2}{9s} \frac{s^2 + u^2}{t^2}. \quad (49.5)$$

If the quarks  $q$  and  $q'$  are identical, we have

$$\frac{d\sigma}{d\Omega}(q\bar{q} \rightarrow q\bar{q}) = \frac{\alpha_s^2}{9s} \left[ \frac{t^2 + u^2}{s^2} + \frac{s^2 + u^2}{t^2} - \frac{2u^2}{3st} \right], \quad (49.6)$$

and by crossing

$$\frac{d\sigma}{d\Omega}(qq \rightarrow qq) = \frac{\alpha_s^2}{9s} \left[ \frac{t^2 + s^2}{u^2} + \frac{s^2 + u^2}{t^2} - \frac{2s^2}{3ut} \right]. \quad (49.7)$$

Annihilation of  $e^+e^-$  into  $\gamma\gamma$  has the cross section

$$\frac{d\sigma}{d\Omega}(e^+e^- \rightarrow \gamma\gamma) = \frac{\alpha^2}{2s} \frac{u^2 + t^2}{tu}. \quad (49.8)$$

The related QCD process also has a triple-gluon coupling. The cross section is

$$\frac{d\sigma}{d\Omega}(q\bar{q} \rightarrow gg) = \frac{8\alpha_s^2}{27s} (t^2 + u^2) \left( \frac{1}{tu} - \frac{9}{4s^2} \right). \quad (49.9)$$

The crossed reactions are

$$\frac{d\sigma}{d\Omega}(qg \rightarrow qg) = \frac{\alpha_s^2}{9s} (s^2 + u^2) \left( -\frac{1}{su} + \frac{9}{4t^2} \right) \quad (49.10)$$

and

$$\frac{d\sigma}{d\Omega}(gg \rightarrow q\bar{q}) = \frac{\alpha_s^2}{24s} (t^2 + u^2) \left( \frac{1}{tu} - \frac{9}{4s^2} \right). \quad (49.11)$$

Finally,

$$\frac{d\sigma}{d\Omega}(gg \rightarrow gg) = \frac{9\alpha_s^2}{8s} \left( 3 - \frac{ut}{s^2} - \frac{su}{t^2} - \frac{st}{u^2} \right). \quad (49.12)$$

Lepton-quark scattering is analogous (neglecting  $Z$  exchange)

$$\frac{d\sigma}{d\Omega}(eq \rightarrow eq) = \frac{\alpha^2}{2s} e_q^2 \frac{s^2 + u^2}{t^2}. \quad (49.13)$$

where  $e_q$  is the charge of the quark. For neutrino scattering with the four-Fermi interaction

$$\frac{d\sigma}{d\Omega}(\nu d \rightarrow \ell^- u) = \frac{G_F^2 s}{4\pi^2}, \quad (49.14)$$

where the Cabibbo angle suppression is ignored. Similarly

$$\frac{d\sigma}{d\Omega}(\nu \bar{u} \rightarrow \ell^- \bar{d}) = \frac{G_F^2 s}{4\pi^2} \frac{(1 + \cos\theta)^2}{4}. \quad (49.15)$$

To obtain the formulae for deep inelastic scattering (presented in more detail in Section 16) we consider quarks of type  $i$  carrying a fraction  $x = Q^2/(2M\nu)$  of the nucleon's energy, where  $\nu = E - E'$  is the energy lost by the lepton in the nucleon rest frame. With  $y = \nu/E$  we have the correspondences

$$1 + \cos\theta \rightarrow 2(1-y),$$

$$d\Omega_{cm} \rightarrow 4\pi f_i(x) dx dy, \quad (49.16)$$

where the latter incorporates the quark distribution,  $f_i(x)$ . In this way we find

$$\begin{aligned} \frac{d\sigma}{dx dy}(eN \rightarrow eX) &= \frac{4\pi\alpha^2 xs}{Q^4} \frac{1}{2} [1 + (1-y)^2] \\ &\times \left[ \frac{4}{9}(u(x) + \bar{u}(x) + \dots) + \frac{1}{9}(d(x) + \bar{d}(x) + \dots) \right] \end{aligned} \quad (49.17)$$

where now  $s = 2ME$  is the cm energy squared for the electron-nucleon collision and we have suppressed contributions from higher mass quarks.

Similarly,

$$\frac{d\sigma}{dx dy}(\nu N \rightarrow \ell^- X) = \frac{G_F^2 xs}{\pi} [(d(x) + \dots) + (1-y)^2(\bar{u}(x) + \dots)] \quad (49.18)$$

and

$$\frac{d\sigma}{dx dy}(\bar{\nu} N \rightarrow \ell^+ X) = \frac{G_F^2 xs}{\pi} [(\bar{d}(x) + \dots) + (1-y)^2(u(x) + \dots)]. \quad (49.19)$$

Quasi-elastic neutrino scattering ( $\nu_\mu n \rightarrow \mu^- p$ ,  $\bar{\nu}_\mu p \rightarrow \mu^+ n$ ) is directly related to the crossed reaction, neutron decay. The formula for the differential cross section is presented, for example, in N.J. Baker *et al.*, Phys. Rev. **D23**, 2499 (1981).

### 49.3. Hadroproduction of heavy quarks

For hadroproduction of heavy quarks  $Q = c, b, t$ , it is important to include mass effects in the formulae. For  $q\bar{q} \rightarrow Q\bar{Q}$ , one has

$$\frac{d\sigma}{d\Omega}(q\bar{q} \rightarrow Q\bar{Q}) = \frac{\alpha_s^2}{9s^3} \sqrt{1 - \frac{4m_Q^2}{s}} \left[ (m_Q^2 - t)^2 + (m_Q^2 - u)^2 + 2m_Q^2 s \right], \quad (49.20)$$

while for  $gg \rightarrow Q\bar{Q}$  one has

$$\begin{aligned} \frac{d\sigma}{d\Omega}(gg \rightarrow Q\bar{Q}) = & \frac{\alpha_s^2}{32s} \sqrt{1 - \frac{4m_Q^2}{s}} \left[ \frac{6}{s^2} (m_Q^2 - t)(m_Q^2 - u) \right. \\ & - \frac{m_Q^2(s - 4m_Q^2)}{3(m_Q^2 - t)(m_Q^2 - u)} \\ & + \frac{4}{3} \frac{(m_Q^2 - t)(m_Q^2 - u) - 2m_Q^2(m_Q^2 + t)}{(m_Q^2 - t)^2} \\ & + \frac{4}{3} \frac{(m_Q^2 - t)(m_Q^2 - u) - 2m_Q^2(m_Q^2 + u)}{(m_Q^2 - u)^2} \\ & - 3 \frac{(m_Q^2 - t)(m_Q^2 - u) + m_Q^2(u - t)}{s(m_Q^2 - t)} \\ & \left. - 3 \frac{(m_Q^2 - t)(m_Q^2 - u) + m_Q^2(t - u)}{s(m_Q^2 - u)} \right]. \quad (49.21) \end{aligned}$$

### 49.4. Production of Weak Gauge Bosons

#### 49.4.1. $W$ and $Z$ resonant production :

Resonant production of a single  $W$  or  $Z$  is governed by the partial widths

$$\Gamma(W \rightarrow \ell_i \bar{\nu}_i) = \frac{\sqrt{2}G_F m_W^3}{12\pi} \quad (49.22)$$

$$\Gamma(W \rightarrow q_i \bar{q}_j) = 3 \frac{\sqrt{2}G_F |V_{ij}|^2 m_W^3}{12\pi} \quad (49.23)$$

$$\begin{aligned} \Gamma(Z \rightarrow f\bar{f}) = & N_c \frac{\sqrt{2}G_F m_Z^3}{6\pi} \\ & \times \left[ (T_3 - Q_f \sin^2 \theta_W)^2 + (Q_f \sin^2 \theta_W)^2 \right] \quad (49.24) \end{aligned}$$

The weak mixing angle is  $\theta_W$ . The CKM matrix elements are indicated by  $V_{ij}$  and  $N_c$  is 3 for  $q\bar{q}$  final states and 1 for leptonic final states.

The full differential cross section for  $f_i \bar{f}_j \rightarrow (W, Z) \rightarrow f'_i \bar{f}'_{j'}$  is given by

$$\begin{aligned} \frac{d\sigma}{d\Omega} = & \frac{N_c^f}{N_c^{i'}} \cdot \frac{1}{256\pi^2 s} \cdot \frac{s^2}{(s - M^2)^2 + s\Gamma^2} \\ & \times \left[ (L^2 + R^2)(L'^2 + R'^2)(1 + \cos^2 \theta) \right. \\ & \left. + (L^2 - R^2)(L'^2 - R'^2)2 \cos \theta \right] \quad (49.25) \end{aligned}$$

where  $M$  is the mass of the  $W$  or  $Z$ . The couplings for the  $W$  are  $L = (8G_F m_W^2 / \sqrt{2})^{1/2} V_{ij} / \sqrt{2}$ ;  $R = 0$  where  $V_{ij}$  is the corresponding CKM matrix element, with an analogous expression for  $L'$  and  $R'$ . For  $Z$ , the couplings are  $L = (8G_F m_Z^2 / \sqrt{2})^{1/2} (T_3 - \sin^2 \theta_W Q)$ ;  $R = -(8G_F m_Z^2 / \sqrt{2})^{1/2} \sin^2 \theta_W Q$ , where  $T_3$  is the weak isospin of the initial left-handed fermion and  $Q$  is the initial fermion's electric charge. The expressions for  $L'$  and  $R'$  are analogous. The color factors  $N_c^{i,f}$  are 3 for initial or final quarks and 1 for initial or final leptons.

#### 49.4.2. Production of pairs of weak gauge bosons :

The cross section for  $f\bar{f} \rightarrow W^+ W^-$  is given in term of the couplings of the left-handed and right-handed fermion  $f$ ,  $\ell = 2(T_3 - Qx_W)$ ,  $r = -2Qx_W$ , where  $T_3$  is the third component of weak isospin for the left-handed  $f$ ,  $Q$  is its electric charge (in units of the proton charge), and  $x_W = \sin^2 \theta_W$ :

$$\begin{aligned} \frac{d\sigma}{dt} = & \frac{2\pi\alpha^2}{N_c s^2} \left\{ \left[ \left( Q + \frac{\ell + r}{4x_W} \frac{s}{s - m_Z^2} \right)^2 + \left( \frac{\ell - r}{4x_W} \frac{s}{s - m_Z^2} \right)^2 \right] A(s, t, u) \right. \\ & + \frac{1}{2x_W} \left( Q + \frac{\ell}{2x_W} \frac{s}{s - m_Z^2} \right) (\Theta(-Q)I(s, t, u) - \Theta(Q)I(s, u, t)) \\ & \left. + \frac{1}{8x_W^2} (\Theta(-Q)E(s, t, u) + \Theta(Q)E(s, u, t)) \right\}, \quad (49.26) \end{aligned}$$

where  $\Theta(x)$  is 1 for  $x > 0$  and 0 for  $x < 0$ , and where

$$\begin{aligned} A(s, t, u) = & \left( \frac{tu}{m_W^4} - 1 \right) \left( \frac{1}{4} - \frac{m_W^2}{s} + 3 \frac{m_W^4}{s^2} \right) + \frac{s}{m_W^2} - 4, \\ I(s, t, u) = & \left( \frac{tu}{m_W^4} - 1 \right) \left( \frac{1}{4} - \frac{m_W^2}{2s} - \frac{m_W^4}{st} \right) + \frac{s}{m_W^2} - 2 + 2 \frac{m_W^2}{t}, \\ E(s, t, u) = & \left( \frac{tu}{m_W^4} - 1 \right) \left( \frac{1}{4} + \frac{m_W^4}{t^2} \right) + \frac{s}{m_W^2}, \quad (49.27) \end{aligned}$$

and  $s, t, u$  are the usual Mandelstam variables with  $s = (p_f + p_{\bar{f}})^2$ ,  $t = (p_f - p_{W^-})^2$ ,  $u = (p_f - p_{W^+})^2$ . The factor  $N_c$  is 3 for quarks and 1 for leptons.

The analogous cross-section for  $q_i \bar{q}_j \rightarrow W^\pm Z^0$  is

$$\begin{aligned} \frac{d\sigma}{dt} = & \frac{\pi\alpha^2 |V_{ij}|^2}{6s^2 x_W^2} \left\{ \left( \frac{1}{s - m_W^2} \right)^2 \left[ \left( \frac{9 - 8x_W}{4} \right) (ut - m_W^2 m_Z^2) \right. \right. \\ & + (8x_W - 6)s(m_W^2 + m_Z^2) \left. \right] \\ & + \left[ \frac{ut - m_W^2 m_Z^2 - s(m_W^2 + m_Z^2)}{s - m_W^2} \right] \left[ \frac{\ell_j}{t} - \frac{\ell_i}{u} \right] \\ & \left. + \frac{ut - m_W^2 m_Z^2}{4(1 - x_W)} \left[ \frac{\ell_j^2}{t^2} + \frac{\ell_i^2}{u^2} \right] + \frac{s(m_W^2 + m_Z^2)}{2(1 - x_W)} \frac{\ell_i \ell_j}{tu} \right\}, \quad (49.28) \end{aligned}$$

where  $\ell_i$  and  $\ell_j$  are the couplings of the left-handed  $q_i$  and  $q_j$  as defined above. The CKM matrix element between  $q_i$  and  $q_j$  is  $V_{ij}$ .

The cross section for  $q_i \bar{q}_i \rightarrow Z^0 Z^0$  is

$$\frac{d\sigma}{dt} = \frac{\pi\alpha^2}{96} \frac{\ell_i^4 + r_i^4}{x_W^2 (1 - x_W^2)^2 s^2} \left[ \frac{t}{u} + \frac{u}{t} + \frac{4m_Z^2 s}{tu} - m_Z^4 \left( \frac{1}{t^2} + \frac{1}{u^2} \right) \right]. \quad (49.29)$$

### 49.5. Production of Higgs Bosons

#### 49.5.1. Resonant Production :

The Higgs boson of the Standard Model can be produced resonantly in the collisions of quarks, leptons,  $W$  or  $Z$  bosons, gluons, or photons. The production cross section is thus controlled by the partial width of the Higgs boson into the entrance channel and its total width. The branching fractions for the Standard Model Higgs boson are shown in Fig. 1 of the "Searches for Higgs bosons" review in the Particle Listings section, as a function of the Higgs boson mass. The partial widths are given by the relations

$$\Gamma(H \rightarrow f\bar{f}) = \frac{G_F m_f^2 m_H N_c}{4\pi\sqrt{2}} \left(1 - 4m_f^2/m_H^2\right)^{3/2}, \quad (49.30)$$

$$\Gamma(H \rightarrow W^+W^-) = \frac{G_F m_H^3 \beta_W}{32\pi\sqrt{2}} \left(4 - 4a_W + 3a_W^2\right), \quad (49.31)$$

$$\Gamma(H \rightarrow ZZ) = \frac{G_F m_H^3 \beta_Z}{64\pi\sqrt{2}} \left(4 - 4a_Z + 3a_Z^2\right), \quad (49.32)$$

where  $N_c$  is 3 for quarks and 1 for leptons and where  $a_W = 1 - \beta_W^2 = 4m_W^2/m_H^2$  and  $a_Z = 1 - \beta_Z^2 = 4m_Z^2/m_H^2$ . The decay to two gluons proceeds through quark loops, with the  $t$  quark dominating [2]. Explicitly,

$$\Gamma(H \rightarrow gg) = \frac{\alpha_s^2 G_F m_H^3}{36\pi^3 \sqrt{2}} \left| \sum_q I(m_q^2/m_H^2) \right|^2, \quad (49.33)$$

where  $I(z)$  is complex for  $z < 1/4$ . For  $z < 2 \times 10^{-3}$ ,  $|I(z)|$  is small so the light quarks contribute negligibly. For  $m_H < 2m_t$ ,  $z > 1/4$  and

$$I(z) = 3 \left[ 2z + 2z(1-4z) \left( \sin^{-1} \frac{1}{2\sqrt{z}} \right)^2 \right], \quad (49.34)$$

which has the limit  $I(z) \rightarrow 1$  as  $z \rightarrow \infty$ .

#### 49.5.2. Higgs Boson Production in $W^*$ and $Z^*$ decay :

The Standard Model Higgs boson can be produced in the decay of a virtual  $W$  or  $Z$  ("Higgsstrahlung") [3,4]: In particular, if  $k$  is the c.m. momentum of the Higgs boson,

$$\sigma(q_i \bar{q}_j \rightarrow WH) = \frac{\pi \alpha^2 |V_{ij}|^2}{36 \sin^4 \theta_W} \frac{2k}{\sqrt{s}} \frac{k^2 + 3m_W^2}{(s - m_W^2)^2} \quad (49.35)$$

$$\sigma(f\bar{f} \rightarrow ZH) = \frac{2\pi \alpha^2 (\ell_f^2 + r_f^2)}{48 N_c \sin^4 \theta_W \cos^4 \theta_W} \frac{2k}{\sqrt{s}} \frac{k^2 + 3m_Z^2}{(s - m_Z^2)^2}, \quad (49.36)$$

where  $\ell$  and  $r$  are defined as above.

#### 49.5.3. $W$ and $Z$ Fusion :

Just as high-energy electrons can be regarded as sources of virtual photon beams, at very high energies they are sources of virtual  $W$  and  $Z$  beams. For Higgs boson production, it is the longitudinal components of the  $W$ s and  $Z$ s that are important [5]. The distribution of longitudinal  $W$ s carrying a fraction  $y$  of the electron's energy is [6]

$$f(y) = \frac{g^2}{16\pi^2} \frac{1-y}{y}, \quad (49.37)$$

where  $g = e/\sin \theta_W$ . In the limit  $s \gg m_H \gg m_W$ , the partial decay rate is  $\Gamma(H \rightarrow W_L W_L) = (g^2/64\pi)(m_H^3/m_W^2)$  and in the equivalent  $W$  approximation [7]

$$\sigma(e^+e^- \rightarrow \bar{\nu}_e \nu_e H) = \frac{1}{16m_W^2} \left( \frac{\alpha}{\sin^2 \theta_W} \right)^3 \times \left[ \left( 1 + \frac{m_H^2}{s} \right) \log \frac{s}{m_H^2} - 2 + 2 \frac{m_H^2}{s} \right]. \quad (49.38)$$

There are significant corrections to this relation when  $m_H$  is not large compared to  $m_W$  [8]. For  $m_H = 150$  GeV, the estimate is too high by 51% for  $\sqrt{s} = 1000$  GeV, 32% too high at  $\sqrt{s} = 2000$  GeV, and 22% too high at  $\sqrt{s} = 4000$  GeV. Fusion of  $ZZ$  to make a Higgs boson can be treated similarly. Identical formulae apply for Higgs production in the collisions of quarks whose charges permit the emission of a  $W^+$  and a  $W^-$ , except that QCD corrections and CKM matrix elements are required. Even in the absence of QCD corrections, the fine-structure constant ought to be evaluated at the scale of the collision, say  $m_W$ . All quarks contribute to the  $ZZ$  fusion process.

### 49.6. Inclusive hadronic reactions

One-particle inclusive cross sections  $E d^3\sigma/d^3p$  for the production of a particle of momentum  $p$  are conveniently expressed in terms of rapidity  $y$  (see above) and the momentum  $p_T$  transverse to the beam direction (in the c.m.):

$$E \frac{d^3\sigma}{d^3p} = \frac{d^3\sigma}{d\phi dy p_T dp_T^2}. \quad (49.39)$$

In appropriate circumstances, the cross section may be decomposed as a partonic cross section multiplied by the probabilities of finding partons of the prescribed momenta:

$$\sigma_{\text{hadronic}} = \sum_{ij} \int dx_1 dx_2 f_i(x_1) f_j(x_2) d\hat{\sigma}_{\text{partonic}}, \quad (49.40)$$

The probability that a parton of type  $i$  carries a fraction of the incident particle's that lies between  $x_1$  and  $x_1 + dx_1$  is  $f_i(x_1)dx_1$  and similarly for partons in the other incident particle. The partonic collision is specified by its c.m. energy squared  $\hat{s} = x_1 x_2 s$  and the momentum transfer squared  $\hat{t}$ . The final hadronic state is more conveniently specified by the rapidities  $y_1, y_2$  of the two jets resulting from the collision and the transverse momentum  $p_T$ . The connection between the differentials is

$$dx_1 dx_2 d\hat{t} = dy_1 dy_2 \frac{\hat{s}}{s} dp_T^2, \quad (49.41)$$

so that

$$\frac{d^3\sigma}{dy_1 dy_2 dp_T^2} = \frac{\hat{s}}{s} \left[ f_i(x_1) f_j(x_2) \frac{d\hat{\sigma}}{d\hat{t}}(\hat{s}, \hat{t}, \hat{u}) + f_i(x_2) f_j(x_1) \frac{d\hat{\sigma}}{d\hat{t}}(\hat{s}, \hat{u}, \hat{t}) \right], \quad (49.42)$$

where we have taken into account the possibility that the incident parton types might arise from either incident particle. The second term should be dropped if the types are identical:  $i = j$ .

### 49.7. Two-photon processes

In the Weizsäcker-Williams picture, a high-energy electron beam is accompanied by a spectrum of virtual photons of energies  $\omega$  and invariant-mass squared  $q^2 = -Q^2$ , for which the photon number density is

$$dn = \frac{\alpha}{\pi} \left[ 1 - \frac{\omega}{E} + \frac{\omega^2}{E^2} - \frac{m_e^2 \omega^2}{Q^2 E^2} \right] \frac{d\omega}{\omega} \frac{dQ^2}{Q^2}, \quad (49.43)$$

where  $E$  is the energy of the electron beam. The cross section for  $e^+e^- \rightarrow e^+e^-X$  is then [9]

$$d\sigma_{e^+e^- \rightarrow e^+e^-X}(s) = dn_1 dn_2 d\sigma_{\gamma\gamma \rightarrow X}(W^2), \quad (49.44)$$

where  $W^2 = m_X^2$ . Integrating from the lower limit  $Q^2 = m_e^2 \frac{\omega_i^2}{E_i(E_i - \omega_i)}$  to a maximum  $Q^2$  gives

$$\sigma_{e^+e^- \rightarrow e^+e^-X}(s) = \frac{\alpha^2}{\pi^2} \int_{z_{th}}^1 \frac{dz}{z} \times \left[ \left( \ln \frac{Q_{max}^2}{zm_e^2} - 1 \right)^2 f(z) + \frac{1}{3} (\ln z)^3 \right] \sigma_{\gamma\gamma \rightarrow X}(zs), \quad (49.45)$$

where

$$f(z) = \left(1 + \frac{1}{2}z\right)^2 \ln(1/z) - \frac{1}{2}(1-z)(3+z). \quad (49.46)$$

The appropriate value of  $Q_{max}^2$  depends on the properties of the produced system  $X$ . For production of hadronic systems,  $Q_{max}^2 \approx m_p^2$ ,

while for lepton-pair production,  $Q^2 \approx W^2$ . For production of a resonance with spin  $J \neq 1$ , we have

$$\sigma_{e^+e^- \rightarrow e^+e^- R}(s) = (2J+1) \frac{8\alpha^2 \Gamma_{R \rightarrow \gamma\gamma}}{m_R^3} \times \left[ f(m_R^2/s) \left( \ln \frac{m_V^2 s}{m_e^2 m_R^2} - 1 \right)^2 - \frac{1}{3} \left( \ln \frac{s}{M_R^2} \right)^3 \right], \quad (49.47)$$

where  $m_V$  is the mass that enters into the form factor for the  $\gamma\gamma \rightarrow R$  transition, typically  $m_\rho$ .

## PART II: PROCESSES BEYOND THE STANDARD MODEL

### 49.8. Production of supersymmetric particles

In supersymmetric (SUSY) theories (see Supersymmetric Particle Searches in this *Review*), every boson has a fermionic superpartner, and every fermion has a bosonic superpartner. The minimal supersymmetric Standard Model (MSSM) is a direct supersymmetrization of the Standard Model (SM), although a second Higgs doublet is needed to avoid triangle anomalies [10]. Under *soft* SUSY breaking, superpartner masses are lifted above the SM particle masses. In weak scale SUSY, the superpartners are invoked to stabilize the weak scale under radiative corrections, so the superpartners are expected to have masses of order the TeV scale.

#### 49.8.1. Gluino and squark production :

The superpartners of gluons are the color octet, spin- $\frac{1}{2}$  gluinos ( $\tilde{g}$ ), while each helicity component of quark flavor has a spin-0 *squark* partner, *e.g.*  $\tilde{q}_L$  and  $\tilde{q}_R$ . Third generation left- and right- squarks are expected to have large mixing, resulting in mass eigenstates  $\tilde{q}_1$  and  $\tilde{q}_2$ , with  $m_{\tilde{q}_1} < m_{\tilde{q}_2}$  (here,  $q$  denotes any of the SM flavors of quarks and  $\tilde{q}_i$  the corresponding flavor and type ( $i = L, R$  or  $1, 2$ ) of squark). Gluino pair production ( $\tilde{g}\tilde{g}$ ) takes place via either glue-glue or quark-antiquark annihilation [11].

The subprocess cross sections are usually presented as differential distributions in the Mandelstam variables  $s$ ,  $t$  and  $u$ . Note that for a  $2 \rightarrow 2$  scattering subprocess  $ab \rightarrow cd$ , the Mandelstam variable  $s = (p_a + p_b)^2 = (p_c + p_d)^2$ , where  $p_a$  is the 4-momentum of particle  $a$ , and so forth. The variable  $t = (p_c - p_a)^2$ , where  $c$  and  $a$  are taken conventionally to be the most similar particles in the subprocess. The variable  $u$  would then be equal to  $(p_d - p_a)^2$ . Note that since  $s$ ,  $t$  and  $u$  are squares of 4-vectors, they are invariants in any inertial reference frame.

Gluino pair production at hadron colliders is described by:

$$\begin{aligned} \frac{d\sigma}{dt}(gg \rightarrow \tilde{g}\tilde{g}) &= \frac{9\pi\alpha_s^2}{4s^2} \left\{ \frac{2(m_g^2 - t)(m_g^2 - u)}{s^2} \right. \\ &+ \frac{(m_g^2 - t)(m_g^2 - u) - 2m_g^2(m_g^2 + t)}{(m_g^2 - t)^2} \\ &+ \frac{(m_g^2 - t)(m_g^2 - u) - 2m_g^2(m_g^2 + u)}{(m_g^2 - u)^2} + \frac{m_g^2(s - 4m_g^2)}{(m_g^2 - t)(m_g^2 - u)} \\ &\left. - \frac{(m_g^2 - t)(m_g^2 - u) + m_g^2(u - t)}{s(m_g^2 - t)} - \frac{(m_g^2 - t)(m_g^2 - u) + m_g^2(t - u)}{s(m_g^2 - u)} \right\}, \quad (49.48) \end{aligned}$$

where  $\alpha_s$  is the strong fine structure constant. Also,

$$\begin{aligned} \frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{g}\tilde{g}) &= \frac{8\pi\alpha_s^2}{9s^2} \left\{ \frac{4}{3} \left( \frac{m_g^2 - t}{m_q^2 - t} \right)^2 + \frac{4}{3} \left( \frac{m_g^2 - u}{m_q^2 - u} \right)^2 \right. \\ &+ \frac{3}{s^2} [(m_g^2 - t)^2 + (m_g^2 - u)^2 + 2m_g^2 s] - 3 \frac{(m_g^2 - t)^2 + m_g^2 s}{s(m_q^2 - t)} \\ &\left. - 3 \frac{(m_g^2 - u)^2 + m_g^2 s}{s(m_q^2 - u)} + \frac{1}{3} \frac{m_g^2 s}{(m_q^2 - t)(m_q^2 - u)} \right\}. \quad (49.49) \end{aligned}$$

Gluinos can also be produced in association with squarks:  $\tilde{g}\tilde{q}_i$  production, where  $\tilde{q}_i$  represents any of the various types (left-, right- or mixed) and flavors of squarks. The subprocess cross section is independent of whether the squark is the right-, left- or mixed type:

$$\begin{aligned} \frac{d\sigma}{dt}(gq \rightarrow \tilde{g}\tilde{q}_i) &= \frac{\pi\alpha_s^2}{24s^2} \left[ \frac{16}{3}(s^2 + (m_{\tilde{q}_i}^2 - u)^2) + \frac{4}{3}s(m_{\tilde{q}_i}^2 - u) \right] \\ &\times \left( (m_g^2 - u)^2 + (m_{\tilde{q}_i}^2 - m_g^2)^2 + \frac{2sm_g^2(m_{\tilde{q}_i}^2 - m_g^2)}{(m_g^2 - t)} \right). \quad (49.50) \end{aligned}$$

There are many different subprocesses for production of squark pairs. Since left- and right- squarks generally have different masses and different decay patterns, we present the differential cross section for each subprocess of  $\tilde{q}_i$  ( $i = L, R$  or  $1, 2$ ) separately. (In early literature, the following formulae were often combined into a single equation which didn't differentiate the various squark types.) The result for  $gg \rightarrow \tilde{q}_i\tilde{q}_i$  is:

$$\begin{aligned} \frac{d\sigma}{dt}(gg \rightarrow \tilde{q}_i\tilde{q}_i) &= \frac{\pi\alpha_s^2}{4s^2} \left\{ \frac{1}{3} \left( \frac{m_q^2 + t}{m_q^2 - t} \right)^2 + \frac{1}{3} \left( \frac{m_q^2 + u}{m_q^2 - u} \right)^2 \right. \\ &+ \frac{3}{32s^2} (8s(4m_q^2 - s) + 4(u - t)^2) + \frac{7}{12} \\ &- \frac{1}{48} \frac{(4m_q^2 - s)^2}{(m_q^2 - t)(m_q^2 - u)} \\ &+ \frac{3}{32} \frac{[(t - u)(4m_q^2 + 4t - s) - 2(m_q^2 - u)(6m_q^2 + 2t - s)]}{s(m_q^2 - t)} \\ &+ \frac{3}{32} \frac{[(u - t)(4m_q^2 + 4u - s) - 2(m_q^2 - t)(6m_q^2 + 2u - s)]}{s(m_q^2 - u)} \\ &\left. + \frac{7}{96} \frac{[4m_q^2 + 4t - s]}{m_q^2 - t} + \frac{7}{96} \frac{[4m_q^2 + 4u - s]}{m_q^2 - u} \right\}, \quad (49.51) \end{aligned}$$

which has an obvious  $u \leftrightarrow t$  symmetry.

For  $q\bar{q} \rightarrow \tilde{q}_i\tilde{q}_i$  with the same initial and final state flavors, we have

$$\begin{aligned} \frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{q}_i\tilde{q}_i) &= \frac{2\pi\alpha_s^2}{9s^2} \left\{ \frac{1}{(t - m_q^2)^2} + \frac{2}{s^2} - \frac{2/3}{s(t - m_q^2)} \right\} \\ &\times [-st - (t - m_{\tilde{q}_i}^2)^2], \quad (49.52) \end{aligned}$$

while if initial and final state flavors are different ( $q\bar{q} \rightarrow \tilde{q}_i'\tilde{q}_i'$ ) we instead have

$$\frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{q}_i'\tilde{q}_i') = \frac{4\pi\alpha_s^2}{9s^4} [-st - (t - m_{\tilde{q}_i'}^2)^2]. \quad (49.53)$$

If the two initial state quarks are of different flavors, then we have

$$\frac{d\sigma}{dt}(q\bar{q}' \rightarrow \tilde{q}_i\tilde{q}_i') = \frac{2\pi\alpha_s^2}{9s^2} \frac{-st - (t - m_{\tilde{q}_i}^2)^2}{(t - m_{\tilde{q}_i'}^2)^2}. \quad (49.54)$$

If the initial quarks are of different flavor and final state squarks are of different type ( $i \neq j$ ) then

$$\frac{d\sigma}{dt}(q\bar{q}' \rightarrow \tilde{q}_i\tilde{q}_j') = \frac{2\pi\alpha_s^2}{9s^2} \frac{m_g^2 s}{(t - m_{\tilde{q}_j'}^2)^2}. \quad (49.55)$$

For same-flavor initial state quarks, but final state unlike-type squarks, we also have

$$\frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{q}_i\tilde{q}_j) = \frac{2\pi\alpha_s^2}{9s^2} \frac{m_g^2 s}{(t - m_{\tilde{q}_j}^2)^2}. \quad (49.56)$$

There also exist cross sections for quark-quark annihilation to squark pairs. For same flavor quark-quark annihilation to same flavor/same type final state squarks,

$$\begin{aligned} \frac{d\sigma}{dt}(qq \rightarrow \tilde{q}_i \tilde{q}_i) &= \\ &= \frac{\pi \alpha_s^2 m_g^2}{9s^2} \left\{ \frac{1}{(t - m_g^2)^2} + \frac{1}{(u - m_g^2)^2} - \frac{2/3}{(t - m_g^2)(u - m_g^2)} \right\}, \end{aligned} \quad (49.57)$$

while if the final type squarks are different ( $i \neq j$ ), we have

$$\begin{aligned} \frac{d\sigma}{dt}(qq \rightarrow \tilde{q}_i \tilde{q}_j) &= \\ &= \frac{2\pi \alpha_s^2}{9s^2} \left\{ \frac{[-st - (t - m_{\tilde{q}_i}^2)(t - m_{\tilde{q}_j}^2)]}{(t - m_g^2)} + \frac{[-su - (u - m_{\tilde{q}_i}^2)(u - m_{\tilde{q}_j}^2)]}{(u - m_g^2)} \right\}. \end{aligned} \quad (49.58)$$

If initial/final state flavors are different, but final state squark types are the same, then

$$\frac{d\sigma}{dt}(qq' \rightarrow \tilde{q}_i \tilde{q}_i') = \frac{2\pi \alpha_s^2}{9s^2} \frac{m_g^2 s}{(t - m_g^2)^2}. \quad (49.59)$$

If initial quark flavors are different and final squark types are different, then

$$\frac{d\sigma}{dt}(qq' \rightarrow \tilde{q}_i \tilde{q}_j') = \frac{2\pi \alpha_s^2}{9s^2} \frac{-st - (t - m_{\tilde{q}_i}^2)(t - m_{\tilde{q}_j}^2)}{(t - m_g^2)^2}. \quad (49.60)$$

#### 49.8.2. Gluino and squark associated production :

In the MSSM, the charged spin- $\frac{1}{2}$  winos and higgsinos mix to make chargino states  $\chi_1^\pm$  and  $\chi_2^\pm$ , with  $m_{\chi_1^\pm} < m_{\chi_2^\pm}$ . The spin- $\frac{1}{2}$  neutral bino, wino and higgsino fields mix to give four neutralino mass eigenstates  $\chi_{1,2,3,4}^0$  ordered according to mass. We sometimes denote the charginos and neutralinos collectively as -inos for notational simplicity

For gluino and squark production in association with charginos and neutralinos [12], the quark-squark-neutralino couplings\* are defined by the interaction Lagrangian terms  $\mathcal{L}_{f\tilde{f}\tilde{\chi}_i^0} = \left[ iA_{\tilde{\chi}_i^0}^f \tilde{f}_L^\dagger \tilde{\chi}_i^0 P_L f + iB_{\tilde{\chi}_i^0}^f \tilde{f}_R^\dagger \tilde{\chi}_i^0 P_R f + \text{h.c.} \right]$ , where  $A_{\tilde{\chi}_i^0}^f$  and  $B_{\tilde{\chi}_i^0}^f$  are coupling constants involving gauge couplings, neutralino mixing elements and in the case of third generation fermions, Yukawa couplings. Their form depends on the conventions used for setting up the MSSM Lagrangian, and can be found in various reviews [13] and textbooks [14,15].  $P_L$  and  $P_R$  are the usual left- and right-spinor projection operators and  $f$  denotes any of the SM fermions  $u, d, e, \nu_e, \dots$ . The fermion-sfermion- chargino couplings have the form  $\mathcal{L} = \left[ iA_{\tilde{\chi}_i^\pm}^d \tilde{u}_L^\dagger \tilde{\chi}_i^\pm P_L d + iA_{\tilde{\chi}_i^\pm}^u \tilde{d}_L^\dagger \tilde{\chi}_i^\pm P_L u + \text{h.c.} \right]$  for  $u$  and  $d$  quarks, where the  $A_{\tilde{\chi}_i^\pm}^d$  and  $A_{\tilde{\chi}_i^\pm}^u$  couplings are again convention-dependent, and can be found in textbooks. The superscript  $c$  denotes “charge conjugate spinor”, defined by  $\psi^c \equiv C\bar{\psi}^T$ .

The subprocess cross sections for chargino-squark associated production occur via squark exchange and are given by

$$\frac{d\sigma}{dt}(\bar{u}g \rightarrow \tilde{\chi}_i^- \tilde{d}_L) = \frac{\alpha_s}{24s^2} |A_{\tilde{\chi}_i^-}^u|^2 \psi(m_{\tilde{d}_L}, m_{\tilde{\chi}_i^-}, t), \quad (49.61)$$

$$\frac{d\sigma}{dt}(dg \rightarrow \tilde{\chi}_i^- \tilde{u}_L) = \frac{\alpha_s}{24s^2} |A_{\tilde{\chi}_i^-}^d|^2 \psi(m_{\tilde{u}_L}, m_{\tilde{\chi}_i^-}, t), \quad (49.62)$$

\* The couplings  $A_{\tilde{\chi}_i^0}^f$  and  $B_{\tilde{\chi}_i^0}^f$  are given explicitly in Ref. 15 in Eq. (8.87). Also, the couplings  $A_{\tilde{\chi}_i^\pm}^d$  and  $A_{\tilde{\chi}_i^\pm}^u$  are given in Eq. (8.93). The couplings  $X_i^j$  and  $Y_i^j$  are given by Eq. (8.103), while the  $x_i$  and  $y_i$  couplings are given in Eq. (8.100). Finally, the couplings  $W_{ij}$  are given in Eq. (8.101).

while neutralino-squark production is given by

$$\frac{d\sigma}{dt}(qg \rightarrow \tilde{\chi}_i^0 \tilde{q}) = \frac{\alpha_s}{24s^2} \left( |A_{\tilde{\chi}_i^0}^q|^2 + |B_{\tilde{\chi}_i^0}^q|^2 \right) \psi(m_{\tilde{q}}, m_{\tilde{\chi}_i^0}, t), \quad (49.63)$$

where

$$\begin{aligned} \psi(m_1, m_2, t) &= \frac{s + t - m_1^2}{2s} - \frac{m_1^2(m_2^2 - t)}{(m_1^2 - t)^2} \\ &+ \frac{t(m_2^2 - m_1^2) + m_2^2(s - m_2^2 + m_1^2)}{s(m_1^2 - t)}. \end{aligned} \quad (49.64)$$

Here, the variable  $t$  is given by the square of “squark-minus-quark” four-momentum. The neutralino-gluino associated production cross section also occurs via squark exchange and is given by

$$\begin{aligned} \frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{\chi}_i^0 \tilde{g}) &= \frac{\alpha_s}{18s^2} \left( |A_{\tilde{\chi}_i^0}^q|^2 + |B_{\tilde{\chi}_i^0}^q|^2 \right) \left[ \frac{(m_{\tilde{\chi}_i^0}^2 - t)(m_g^2 - t)}{(m_g^2 - t)^2} \right. \\ &+ \left. \frac{(m_{\tilde{\chi}_i^0}^2 - u)(m_g^2 - u)}{(m_g^2 - u)^2} - \frac{2\eta_i \eta_{\tilde{g}} m_{\tilde{g}} m_{\tilde{\chi}_i^0} s}{(m_g^2 - t)(m_g^2 - u)} \right], \end{aligned} \quad (49.65)$$

where  $\eta_i$  is the sign of the neutralino mass eigenvalue and  $\eta_{\tilde{g}}$  is the sign of the gluino mass eigenvalue. We also have chargino-gluino associated production:

$$\begin{aligned} \frac{d\sigma}{dt}(\bar{u}d \rightarrow \tilde{\chi}_i^- \tilde{g}) &= \frac{\alpha_s}{18s^2} \left[ |A_{\tilde{\chi}_i^-}^u|^2 \frac{(m_{\tilde{\chi}_i^-}^2 - t)(m_g^2 - t)}{(m_{\tilde{d}_L}^2 - t)^2} \right. \\ &+ |A_{\tilde{\chi}_i^-}^d|^2 \frac{(m_{\tilde{\chi}_i^-}^2 - u)(m_g^2 - u)}{(m_{\tilde{u}_L}^2 - u)^2} + \left. \frac{2\eta_{\tilde{g}} \text{Re}(A_{\tilde{\chi}_i^-}^u A_{\tilde{\chi}_i^-}^d) m_{\tilde{g}} m_{\tilde{\chi}_i^-} s}{(m_{\tilde{d}_L}^2 - t)(m_{\tilde{u}_L}^2 - u)} \right], \end{aligned} \quad (49.66)$$

where  $\hat{t} = (\tilde{g} - d)^2$  and in the third term one must take the real part of the in general complex coupling constant product.

#### 49.8.3. Slepton and sneutrino production :

The subprocess cross section for  $\tilde{\ell}_L \tilde{\nu}_{\ell_L}$  production ( $\ell = e$  or  $\mu$ ) occurs via  $s$ -channel  $W$  exchange and is given by

$$\frac{d\sigma}{dt}(d\bar{u} \rightarrow \tilde{\ell}_L \tilde{\nu}_{\ell_L}) = \frac{g^4 |D_W(s)|^2}{192\pi s^2} \left( tu - m_{\tilde{\ell}_L}^2 m_{\tilde{\nu}_{\ell_L}}^2 \right), \quad (49.67)$$

where  $D_W(s) = 1/(s - M_W^2 + iM_W\Gamma_W)$  is the  $W$ -boson propagator denominator. The production of  $\tilde{\tau}_1 \tilde{\nu}_\tau$  is given as above, but replacing  $m_{\tilde{\ell}_L} \rightarrow m_{\tilde{\tau}_1}$ ,  $m_{\tilde{\nu}_{\ell_L}} \rightarrow m_{\tilde{\nu}_\tau}$  and multiplying by an overall factor of  $\cos^2 \theta_\tau$  (where  $\theta_\tau$  is the tau-slepton mixing angle). Similar substitutions hold for  $\tilde{\tau}_2 \tilde{\nu}_\tau$  production, except the overall factor is  $\sin^2 \theta_\tau$ .

**Table 49.1:** The constants  $\alpha_f$  and  $\beta_f$  that appear in in the SM neutral current Lagrangian. Here  $t \equiv \tan \theta_W$  and  $c \equiv \cot \theta_W$ .

$f$	$q_f$	$\alpha_f$	$\beta_f$
$\ell$	-1	$\frac{1}{4}(3t - c)$	$\frac{1}{4}(t + c)$
$\nu_\ell$	0	$\frac{1}{4}(t + c)$	$-\frac{1}{4}(t + c)$
$u$	$\frac{2}{3}$	$-\frac{5}{12}t + \frac{1}{4}c$	$-\frac{1}{4}(t + c)$
$d$	$-\frac{1}{3}$	$\frac{1}{12}t - \frac{1}{4}c$	$\frac{1}{4}(t + c)$

The subprocess cross section for  $\tilde{\ell}_L \bar{\ell}_L$  production occurs via  $s$ -channel  $\gamma$  and  $Z$  exchange, and depends on the neutral current interaction, with fermion couplings to  $\gamma$  and  $Z^0$  given by  $\mathcal{L}_{\text{neutral}} = -e q_f \bar{f} \gamma^\mu f A_\mu + e \bar{f} \gamma^\mu (\alpha_f + \beta_f \gamma_5) f Z_\mu$  (with values of  $q_f$ ,  $\alpha_f$ , and  $\beta_f$  given in Table 49.1).

The subprocess cross section is given by

$$\frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{\ell}_L \bar{\ell}_L) = \frac{e^4}{24\pi s^2} (tu - m_{\tilde{\ell}_L}^4) \times \left\{ \frac{q_\ell^2 q_q^2}{s^2} + (\alpha_\ell - \beta_\ell)^2 (\alpha_q^2 + \beta_q^2) |D_Z(s)|^2 + \frac{2q_\ell q_q \alpha_q (\alpha_\ell - \beta_\ell) (s - M_Z^2)}{s} |D_Z(s)|^2 \right\}, \quad (49.68)$$

where  $D_Z(s) = 1/(s - M_Z^2 + iM_Z\Gamma_Z)$ . The cross section for sneutrino production is given by the same formula, but with  $\alpha_\ell$ ,  $\beta_\ell$ ,  $q_\ell$  and  $m_{\tilde{\ell}_L}$  replaced by  $\alpha_\nu$ ,  $\beta_\nu$ , 0 and  $m_{\tilde{\nu}_L}$ , respectively. The cross section for  $\tilde{\tau}_1 \bar{\tau}_1$  production is obtained by replacing  $m_{\tilde{\ell}_L} \rightarrow m_{\tilde{\tau}_1}$  and  $\beta_\ell \rightarrow \beta_\ell \cos 2\theta_\tau$ .

The cross section for  $\tilde{\ell}_R \bar{\ell}_R$  production is given by substituting  $\alpha_\ell - \beta_\ell \rightarrow \alpha_\ell + \beta_\ell$  and  $m_{\tilde{\ell}_L} \rightarrow m_{\tilde{\ell}_R}$  in the equation above. The cross section for  $\tilde{\tau}_2 \bar{\tau}_2$  production is obtained from the formula for  $\tilde{\ell}_R \bar{\ell}_R$  production by replacing  $m_{\tilde{\ell}_R} \rightarrow m_{\tilde{\tau}_2}$  and  $\beta_\ell \rightarrow \beta_\ell \cos 2\theta_\tau$ .

Finally, the cross section for  $\tilde{\tau}_1 \bar{\tau}_2$  production occurs only via  $Z$  exchange, and is given by

$$\frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{\tau}_1 \bar{\tau}_2) = \frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{\tau}_1 \bar{\tau}_1) = \frac{e^4}{24\pi s^2} (\alpha_q^2 + \beta_q^2) \beta_\ell^2 \sin^2 2\theta_\tau |D_Z(s)|^2 (ut - m_{\tilde{\tau}_1}^2 m_{\tilde{\tau}_2}^2). \quad (49.69)$$

#### 49.8.4. Chargino and neutralino pair production :

##### 49.8.4.1. $\tilde{\chi}_i^- \tilde{\chi}_j^0$ production:

The subprocess cross section for  $d\bar{u} \rightarrow \tilde{\chi}_i^- \tilde{\chi}_j^0$  depends on Lagrangian couplings  $\mathcal{L}_{W\bar{u}d} = -\frac{g}{\sqrt{2}} \bar{u} \gamma_\mu P_L d W^{+\mu} + \text{h.c.}$ ,  $\mathcal{L}_{W\tilde{\chi}_i^- \tilde{\chi}_j^0} = -g(-i)^{\theta_j} \bar{\tilde{\chi}}_i^- [X_i^j + Y_i^j \gamma_5] \gamma_\mu \tilde{\chi}_j^0 W^{-\mu} + \text{h.c.}$ ,  $\mathcal{L}_{q\bar{q}\tilde{\chi}_i^-} = iA_{\tilde{\chi}_i^-}^d \bar{u}_L \tilde{\chi}_i^- P_L d + iA_{\tilde{\chi}_i^-}^u \bar{d}_L \tilde{\chi}_i^- P_L u + \text{h.c.}$  and  $\mathcal{L}_{q\bar{q}\tilde{\chi}_j^0} = iA_{\tilde{\chi}_j^0}^q \bar{q}_L \tilde{\chi}_j^0 P_L q + \text{h.c.}$ . Contributing diagrams include  $W$  exchange and also  $\tilde{d}_L$  and  $\tilde{u}_L$  squark exchange. The  $X_i^j$  and  $Y_i^j$  couplings are new, and again convention-dependent: the cross section formulae works if the interaction Lagrangian is written in the above form, so that the couplings can be suitably extracted. The term  $\theta_j = 0$  (1) if  $m_{\tilde{\chi}_j^0} > 0$  ( $< 0$ ); it comes about because the neutralino field must be re-defined by a  $-i\gamma_5$  transformation if its mass eigenvalue is negative [15]. The subprocess cross section is given in terms of dot products of four momenta, where particle labels are used to denote their four-momenta; note that all mass terms in the cross section formulae are positive definite, so that the signs of mass eigenstates have been absorbed into the Lagrangian couplings, as for instance in Ref. [15]. We then have

$$\frac{d\sigma}{dt}(d\bar{u} \rightarrow \tilde{\chi}_i^- \tilde{\chi}_j^0) = \frac{1}{192\pi s^2} \left[ T_W + T_{\tilde{d}_L} + T_{\tilde{u}_L} + T_{W\tilde{d}_L} + T_{W\tilde{u}_L} + T_{\tilde{d}_L \tilde{u}_L} \right] \quad (49.70)$$

where

$$T_W = 8g^4 |D_W(s)|^2 \left\{ [X_i^{j2} + Y_i^{j2}] (\tilde{\chi}_j^0 \cdot d \tilde{\chi}_i^- \cdot \bar{u} + \tilde{\chi}_j^0 \cdot \bar{u} \tilde{\chi}_i^- \cdot d) + 2(X_i^j Y_i^j) (\tilde{\chi}_j^0 \cdot d \tilde{\chi}_i^- \cdot \bar{u} - \tilde{\chi}_j^0 \cdot \bar{u} \tilde{\chi}_i^- \cdot d) + [X_i^{j2} - Y_i^{j2}] m_{\tilde{\chi}_i^-} m_{\tilde{\chi}_j^0} d \cdot \bar{u} \right\}, \quad (49.71)$$

$$T_{\tilde{d}_L} = \frac{4|A_{\tilde{\chi}_i^-}^d|^2 |A_{\tilde{\chi}_j^0}^d|^2}{[(\tilde{\chi}_i^- - \bar{u})^2 - m_{\tilde{d}_L}^2]^2} d \cdot \tilde{\chi}_j^0 \tilde{\chi}_i^- \cdot \bar{u}, \quad (49.72)$$

$$T_{\tilde{u}_L} = \frac{4|A_{\tilde{\chi}_i^-}^d|^2 |A_{\tilde{\chi}_j^0}^u|^2}{[(\tilde{\chi}_j^0 - \bar{u})^2 - m_{\tilde{u}_L}^2]^2} \bar{u} \cdot \tilde{\chi}_j^0 \tilde{\chi}_i^- \cdot d \quad (49.73)$$

$$T_{W\tilde{d}_L} = \frac{-\sqrt{2}g^2 \text{Re}[A_{\tilde{\chi}_j^0}^{d*} A_{\tilde{\chi}_i^-}^u (-i)^{\theta_j}] (s - M_W^2) |D_W(s)|^2}{(\tilde{\chi}_i^- - \bar{u})^2 - m_{\tilde{d}_L}^2} \times \left\{ 8(X_i^j + Y_i^j) \tilde{\chi}_j^0 \cdot d \bar{u} \cdot \tilde{\chi}_i^- + 4(X_i^j - Y_i^j) m_{\tilde{\chi}_i^-} m_{\tilde{\chi}_j^0} d \cdot \bar{u} \right\} \quad (49.74)$$

$$T_{W\tilde{u}_L} = \frac{\sqrt{2}g^2 \text{Re}[A_{\tilde{\chi}_j^0}^{d*} A_{\tilde{\chi}_i^-}^u (-i)^{\theta_j}] (s - M_W^2) |D_W(s)|^2}{(\tilde{\chi}_j^0 - \bar{u})^2 - m_{\tilde{u}_L}^2} \times \left\{ 8(X_i^j - Y_i^j) \tilde{\chi}_j^0 \cdot d \bar{u} \cdot \tilde{\chi}_i^- + 4(X_i^j + Y_i^j) m_{\tilde{\chi}_i^-} m_{\tilde{\chi}_j^0} d \cdot \bar{u} \right\} \quad (49.75)$$

and

$$T_{\tilde{d}_L \tilde{u}_L} = -\frac{4 \text{Re}[A_{\tilde{\chi}_j^0}^d A_{\tilde{\chi}_i^-}^{u*} A_{\tilde{\chi}_i^-}^{d*} A_{\tilde{\chi}_j^0}^u] m_{\tilde{\chi}_i^-} m_{\tilde{\chi}_j^0} d \cdot \bar{u}}{[(\tilde{\chi}_i^- - \bar{u})^2 - m_{\tilde{d}_L}^2][(\tilde{\chi}_j^0 - \bar{u})^2 - m_{\tilde{u}_L}^2]}. \quad (49.76)$$

##### 49.8.4.2. Chargino pair production:

The subprocess cross section for  $d\bar{d} \rightarrow \tilde{\chi}_i^- \tilde{\chi}_i^+$  ( $i = 1, 2$ ) depends on Lagrangian couplings  $\mathcal{L} = e \tilde{\chi}_i^- \gamma_\mu \tilde{\chi}_i^+ A^\mu - e \cot \theta_W \tilde{\chi}_i^- \gamma_\mu (x_i - y_i \gamma_5) \tilde{\chi}_i^+ Z^\mu$  and also  $\mathcal{L} \ni iA_{\tilde{\chi}_i^-}^d \bar{u}_L \tilde{\chi}_i^- P_L d + iA_{\tilde{\chi}_i^-}^u \bar{d}_L \tilde{\chi}_i^- P_L u + \text{h.c.}$ . Contributing diagrams include  $s$ -channel  $\gamma$ ,  $Z^0$  exchange and  $t$ -channel  $\tilde{u}_L$  exchange [16,17]. The couplings  $x_i$  and  $y_i$  are again new and as usual convention-dependent.

The subprocess cross section is given by

$$\frac{d\sigma}{dt}(d\bar{d} \rightarrow \tilde{\chi}_i^- \tilde{\chi}_i^+) = \frac{1}{192\pi s^2} [T_\gamma + T_Z + T_{\tilde{u}_L} + T_{\gamma Z} + T_{\gamma \tilde{u}_L} + T_{Z \tilde{u}_L}] \quad (49.77)$$

where

$$T_\gamma = \frac{32e^4 q_d^2}{s^2} \left[ d \cdot \tilde{\chi}_i^+ \bar{d} \cdot \tilde{\chi}_i^- + d \cdot \tilde{\chi}_i^- \bar{d} \cdot \tilde{\chi}_i^+ + m_{\tilde{\chi}_i^-}^2 d \cdot \bar{d} \right] \quad (49.78)$$

$$T_Z = 32e^4 \cot^2 \theta_W |D_Z(s)|^2$$

$$\left\{ (\alpha_d^2 + \beta_d^2) (x_i^2 + y_i^2) \left[ d \cdot \tilde{\chi}_i^+ \bar{d} \cdot \tilde{\chi}_i^- + d \cdot \tilde{\chi}_i^- \bar{d} \cdot \tilde{\chi}_i^+ + m_{\tilde{\chi}_i^-}^2 d \cdot \bar{d} \right] \mp 4\alpha_d \beta_d x_i y_i [d \cdot \tilde{\chi}_i^+ \bar{d} \cdot \tilde{\chi}_i^- - d \cdot \tilde{\chi}_i^- \bar{d} \cdot \tilde{\chi}_i^+] - 2y_i^2 (\alpha_d^2 + \beta_d^2) m_{\tilde{\chi}_i^-}^2 d \cdot \bar{d} \right\}, \quad (49.79)$$

$$T_{\tilde{u}_L} = \frac{4|A_{\tilde{\chi}_i^-}^d|^4}{[(d - \tilde{\chi}_i^-)^2 - m_{\tilde{u}_L}^2]^2} d \cdot \tilde{\chi}_i^- \bar{d} \cdot \tilde{\chi}_i^+ \quad (49.80)$$

$$T_{\gamma Z} = \frac{64e^4 \cot \theta_W q_d (s - M_Z^2) |D_Z(s)|^2}{s} \times$$

$$\left\{ \alpha_d x_i \left( d \cdot \tilde{\chi}_i^+ \bar{d} \cdot \tilde{\chi}_i^- + d \cdot \tilde{\chi}_i^- \bar{d} \cdot \tilde{\chi}_i^+ + m_{\tilde{\chi}_i^-}^2 d \cdot \bar{d} \right) \pm \beta_d y_i (d \cdot \tilde{\chi}_i^- \bar{d} \cdot \tilde{\chi}_i^+ - d \cdot \tilde{\chi}_i^+ \bar{d} \cdot \tilde{\chi}_i^-) \right\} \quad (49.81)$$

$$T_{\gamma \tilde{u}_L} = \mp \frac{8e^2 q_d}{s} \frac{|A_{\tilde{\chi}_i^-}^d|^2}{[(d - \tilde{\chi}_i^-)^2 - m_{\tilde{u}_L}^2]} \left\{ 2\bar{d} \cdot \tilde{\chi}_i^+ d \cdot \tilde{\chi}_i^- + m_{\tilde{\chi}_i^-}^2 d \cdot \bar{d} \right\} \quad (49.82)$$

and

$$T_{Z\tilde{u}_L} = \mp 8e^2 \cot \theta_W |D_Z(s)|^2 \frac{|A_{\tilde{\chi}_i^-}^d|^2 (s - M_Z^2)}{[(d - \tilde{\chi}_i^-)^2 - m_{\tilde{u}_L}^2]} (\alpha_d - \beta_d) \\ \times \left\{ 2(x_i \mp y_i) d \cdot \tilde{\chi}_i^- \tilde{d} \cdot \tilde{\chi}_i^+ + m_{\tilde{\chi}_i^-}^2 (x_i \pm y_i) d \cdot \tilde{d} \right\} \quad (49.83)$$

using the upper of the sign choices.

The cross section for  $u\bar{u} \rightarrow \tilde{\chi}_i^+ \tilde{\chi}_i^-$  can be obtained from the above by replacing  $\alpha_d \rightarrow \alpha_u$ ,  $\beta_d \rightarrow \beta_u$ ,  $q_d \rightarrow q_u$ ,  $\tilde{u}_L \rightarrow \tilde{d}_L$ ,  $A_{\tilde{\chi}_i^-}^d \rightarrow A_{\tilde{\chi}_i^-}^u$ ,  $d \rightarrow \bar{u}$ ,  $\bar{d} \rightarrow u$  and adopting the lower of the sign choices everywhere.

The cross section for  $q\bar{q} \rightarrow \tilde{\chi}_1^- \tilde{\chi}_2^+$ ,  $\tilde{\chi}_1^+ \tilde{\chi}_2^-$  can occur via  $Z$  and  $\tilde{q}_L$  exchange. It is usually much smaller than  $\tilde{\chi}_{1,2}^- \tilde{\chi}_{1,2}^+$  production, so the cross section will not be presented here. It can be found in Appendix A of Ref. 15.

#### 49.8.4.3. Neutralino pair production:

Neutralino pair production via  $q\bar{q}$  fusion takes place via  $s$ -channel  $Z$  exchange plus  $t$ - and  $u$ -channel left- and right- squark exchange (5 diagrams) [17,18]. The Lagrangian couplings (see previous footnote\*) needed include terms given above plus terms of the form  $\mathcal{L} = W_{ij} \tilde{\chi}_i^0 \gamma_\mu (\gamma_5)^{\theta_i + \theta_j + 1} \tilde{\chi}_j^0 Z^\mu$ . The couplings  $W_{ij}$  depend only on the *higgsino* components of the neutralinos  $i$  and  $j$ . The subprocess cross section is given by:

$$\frac{d\sigma}{dt}(q\bar{q} \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_j^0) = \frac{1}{192\pi s^2} [T_Z + T_{\tilde{q}_L} + T_{\tilde{q}_R} + T_{Z\tilde{q}_L} + T_{Z\tilde{q}_R}] \quad (49.84)$$

where

$$T_Z = 128e^2 |W_{ij}|^2 (\alpha_q^2 + \beta_q^2) |D_Z(s)|^2 \\ \left[ q \cdot \tilde{\chi}_i^0 \tilde{q} \cdot \tilde{\chi}_j^0 + q \cdot \tilde{\chi}_j^0 \tilde{q} \cdot \tilde{\chi}_i^0 - \eta_i \eta_j m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q} \right], \quad (49.85)$$

$$T_{\tilde{q}_L} = 4|A_{\tilde{\chi}_i^0}^q|^2 |A_{\tilde{\chi}_j^0}^q|^2 \left\{ \frac{q \cdot \tilde{\chi}_i^0 \tilde{q} \cdot \tilde{\chi}_j^0}{[(\tilde{\chi}_i^0 - q)^2 - m_{\tilde{q}_L}^2]^2} + \frac{q \cdot \tilde{\chi}_j^0 \tilde{q} \cdot \tilde{\chi}_i^0}{[(\tilde{\chi}_j^0 - q)^2 - m_{\tilde{q}_L}^2]^2} \right. \\ \left. - \eta_i \eta_j \frac{m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q}}{[(\tilde{\chi}_i^0 - q)^2 - m_{\tilde{q}_L}^2][(\tilde{\chi}_j^0 - q)^2 - m_{\tilde{q}_L}^2]} \right\} \quad (49.86)$$

$$T_{\tilde{q}_R} = 4|B_{\tilde{\chi}_i^0}^q|^2 |B_{\tilde{\chi}_j^0}^q|^2 \left\{ \frac{q \cdot \tilde{\chi}_i^0 \tilde{q} \cdot \tilde{\chi}_j^0}{[(\tilde{\chi}_i^0 - q)^2 - m_{\tilde{q}_R}^2]^2} + \frac{q \cdot \tilde{\chi}_j^0 \tilde{q} \cdot \tilde{\chi}_i^0}{[(\tilde{\chi}_j^0 - q)^2 - m_{\tilde{q}_R}^2]^2} \right. \\ \left. - \eta_i \eta_j \frac{m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q}}{[(\tilde{\chi}_i^0 - q)^2 - m_{\tilde{q}_R}^2][(\tilde{\chi}_j^0 - q)^2 - m_{\tilde{q}_R}^2]} \right\} \quad (49.87)$$

$$T_{Z\tilde{q}_L} = 16e(\alpha_q - \beta_q)(s - M_Z^2) |D_Z(s)|^2 \\ \left\{ \frac{\text{Re}(W_{ij} A_{\tilde{\chi}_i^0}^{q*} A_{\tilde{\chi}_j^0}^q)}{[(\tilde{\chi}_i^0 - q)^2 - m_{\tilde{q}_L}^2]} \left[ 2q \cdot \tilde{\chi}_i^0 \tilde{q} \cdot \tilde{\chi}_j^0 - \eta_i \eta_j m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q} \right] \right. \\ \left. + \eta_i \eta_j \frac{\text{Re}(W_{ij} A_{\tilde{\chi}_i^0}^q A_{\tilde{\chi}_j^0}^{q*})}{[(\tilde{\chi}_j^0 - q)^2 - m_{\tilde{q}_L}^2]} \left[ 2q \cdot \tilde{\chi}_j^0 \tilde{q} \cdot \tilde{\chi}_i^0 - \eta_i \eta_j m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q} \right] \right\} \quad (49.88)$$

$$T_{Z\tilde{q}_R} = 16e(\alpha_q + \beta_q)(s - M_Z^2) |D_Z(s)|^2 \\ \left\{ \frac{\text{Re}(W_{ij} B_{\tilde{\chi}_i^0}^{q*} B_{\tilde{\chi}_j^0}^q)}{[(\tilde{\chi}_i^0 - q)^2 - m_{\tilde{q}_R}^2]} \left[ 2q \cdot \tilde{\chi}_i^0 \tilde{q} \cdot \tilde{\chi}_j^0 - \eta_i \eta_j m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q} \right] \right. \\ \left. - \eta_i \eta_j \frac{\text{Re}(W_{ij} B_{\tilde{\chi}_i^0}^q B_{\tilde{\chi}_j^0}^{q*})}{[(\tilde{\chi}_j^0 - q)^2 - m_{\tilde{q}_R}^2]} \left[ 2q \cdot \tilde{\chi}_j^0 \tilde{q} \cdot \tilde{\chi}_i^0 - \eta_i \eta_j m_{\tilde{\chi}_i^0} m_{\tilde{\chi}_j^0} q \cdot \tilde{q} \right] \right\}. \quad (49.89)$$

As before,  $\eta_i = \pm 1$  corresponding to whether the neutralino mass eigenvalue is positive or negative. When  $i = j$  in the above formula, one must remember to integrate over just  $2\pi$  steradians of solid angle to avoid double counting in the total cross section.

## 49.9. Universal extra dimensions

In the Universal Extra Dimension (UED) model of Ref. [19] (see Ref. [20] for a review of models with extra spacetime dimensions), the Standard Model is embedded in a five dimensional theory, where the fifth dimension is compactified on an  $S_1/Z_2$  orbifold. Each SM chirality state is then the zero mode of an infinite tower of Kaluza-Klein excitations labelled by  $n = 0 - \infty$ . A KK parity is usually assumed to hold, where each state is assigned KK-parity  $P = (-1)^n$ . If the compactification scale is around a TeV, then the  $n = 1$  (or even higher) KK modes may be accessible to collider searches.

Of interest for hadron colliders are the production of massive  $n \geq 1$  quark or gluon pairs. These production cross sections have been calculated in Ref. [21,22]. We list here results for the  $n = 1$  case only with  $M_1 = 1/R$  ( $R$  is the compactification radius) and  $s$ ,  $t$  and  $u$  are the usual Mandelstam variables; more general formulae can be found in Ref. [22]. The superscript  $*$  stands for any KK excited state, while  $\bullet$  stands for left chirality states and  $\circ$  stands for right chirality states.

$$\frac{d\sigma}{dt} = \frac{1}{16\pi s^2} T \quad (49.90)$$

where

$$T(q\bar{q} \rightarrow g^* g^*) = \frac{2g_s^4}{27} \left[ M_1^2 \left( -\frac{4s^3}{t'^2 u'^2} + \frac{57s}{t' u'} - \frac{108}{s} \right) \right. \\ \left. + \frac{20s^2}{t' u'} - 93 + \frac{108t' u'}{s^2} \right] \quad (49.91)$$

and

$$T(gg \rightarrow g^* g^*) = \frac{9g_s^4}{27} \left[ 3M_1^4 \frac{s^2 + t'^2 + u'^2}{t'^2 u'^2} - 3M_1^2 \frac{s^2 + t'^2 + u'^2}{st' u'} + 1 \right. \\ \left. + \frac{(s^2 + t'^2 + u'^2)^3}{4s^2 t'^2 u'^2} - \frac{t' u'}{s^2} \right] \quad (49.92)$$

where  $t' = t - M_1^2$  and  $u' = u - M_1^2$ .

Also,

$$T(q\bar{q} \rightarrow q_1^* \bar{q}_1^*) = \frac{4g_s^4}{9} \left[ \frac{2M_1^2}{s} + \frac{t'^2 + u'^2}{s^2} \right],$$

$$T(q\bar{q} \rightarrow q_1^* \bar{q}_1^*) = \frac{g_s^4}{9} \left[ 2M_1^2 \left( \frac{4}{s} + \frac{s}{t'^2} - \frac{1}{t'} \right) \right. \\ \left. + \frac{23}{6} + \frac{2s^2}{t'^2} + \frac{8s}{3t'} + \frac{6t'}{s} + \frac{8t'^2}{s^2} \right],$$

$$T(qq \rightarrow q_1^* \bar{q}_1^*) = \frac{g_s^4}{27} \left[ M_1^2 \left( 6 \frac{t'}{u'^2} + 6 \frac{u'}{t'^2} - \frac{s}{t' u'} \right) \right. \\ \left. + 2 \left( 3 \frac{t'^2}{u'^2} + 3 \frac{u'^2}{t'^2} + 4 \frac{s^2}{t' u'} - 5 \right) \right],$$

$$T(gg \rightarrow q_1^* \bar{q}_1^*) = g_s^4 \left[ M_1^4 \frac{-4}{t' u'} \left( \frac{s^2}{6t' u'} - \frac{3}{8} \right) \right. \\ \left. + M_1^2 \frac{4}{s} \left( \frac{s^2}{6t' u'} - \frac{3}{8} \right) + \frac{s^2}{6t' u'} - \frac{17}{24} + \frac{3t' u'}{4s^2} \right],$$

$$T(gq \rightarrow g^* q_1^*) = \frac{-g_s^4}{3} \left[ \frac{5s^2}{12t'^2} + \frac{s^3}{t'^2 u'} + \frac{11su'}{6t'^2} + \frac{5u'^2}{12t'^2} + \frac{u'^3}{st'^2} \right],$$

$$T(q\bar{q}' \rightarrow q_1^* \bar{q}_1^*) = \frac{g_s^4}{18} \left[ 4M_1^4 \frac{s}{t'^2} + 5 + 4 \frac{s^2}{t'^2} + 8 \frac{s}{t'} \right],$$

$$T(qq' \rightarrow q_1^* \bar{q}_1^*) = \frac{2g_s^4}{9} \left[ -M_1^2 \frac{s}{t'^2} + \frac{1}{4} + \frac{s^2}{t'^2} \right],$$



$$T(qq \rightarrow q_1^\bullet q_1^\circ) = \frac{g_s^4}{9} \left[ M_1^2 \left( \frac{2s^3}{t'^2 u'^2} - \frac{4s}{t' u'} \right) + 2 \frac{s^4}{t'^2 u'^2} - 8 \frac{s^2}{t' u'} + 5 \right],$$

$$T(q\bar{q}' \rightarrow q_1^\bullet \bar{q}_1^\circ) = \frac{g_s^4}{9} \left[ 2M_1^2 \left( \frac{1}{t'} + \frac{u'}{t'^2} \right) + \frac{5}{2} + \frac{4u'}{t'} + \frac{2u'^2}{t'^2} \right],$$

and

$$T(qq' \rightarrow q_1^\bullet q_1^\circ) = \frac{g_s^4}{9} \left[ -2M_1^2 \left( \frac{1}{t'} + \frac{u'}{t'^2} \right) + \frac{1}{2} + \frac{2u'^2}{t'^2} \right].$$

#### 49.10. Large extra dimensions

In the ADD theory [23] with large extra dimensions (LED), the SM particles are confined to a 3-brane, while gravity propagates in the bulk. It is assumed that the  $n$  extra dimensions are compactified on an  $n$ -dimensional torus of volume  $(2\pi r)^n$ , so that the fundamental  $4+n$  dimensional Planck scale  $M_*$  is related to the usual 4-dimensional Planck scale  $M_{Pl}$  by  $M_{Pl}^2 = M_*^{n+2} (2\pi r)^n$ . If  $M_* \sim 1$  TeV, then the  $M_W - M_{Pl}$  hierarchy problem is just due to gravity propagating in the large extra dimensions.

In these theories, the KK-excited graviton states  $G_{\mu\nu}^n$  for  $n = 1 - \infty$  can be produced at collider experiments. The graviton couplings to matter are suppressed by  $1/M_{Pl}$ , so that graviton emission cross sections  $d\sigma/dt \sim 1/M_{Pl}^2$ . However, the mass splittings between the excited graviton states can be tiny, so the graviton eigenstates are usually approximated by a continuum distribution. A summation (integration) over all allowed graviton emissions ends up cancelling the  $1/M_{Pl}^2$  factor, so that observable cross section rates can be attained. Some of the fundamental production formulae for a KK graviton (denoted  $G$ ) of mass  $m$  at hadron colliders include the subprocesses

$$\frac{d\sigma_m}{dt}(f\bar{f} \rightarrow \gamma G) = \frac{\alpha Q_f^2}{16N_f s M_{Pl}^2} F_1\left(\frac{t}{s}, \frac{m^2}{s}\right), \quad (49.93)$$

where  $Q_f$  is the charge of fermion  $f$  and  $N_f$  is the number of QCD colors of  $f$ . Also,

$$\frac{d\sigma_m}{dt}(q\bar{q} \rightarrow gG) = \frac{\alpha_s}{36 s M_{Pl}^2} F_1\left(\frac{t}{s}, \frac{m^2}{s}\right), \quad (49.94)$$

$$\frac{d\sigma_m}{dt}(qg \rightarrow qG) = \frac{\alpha_s}{96 s M_{Pl}^2} F_2\left(\frac{t}{s}, \frac{m^2}{s}\right), \quad (49.95)$$

$$\frac{d\sigma_m}{dt}(gg \rightarrow gG) = \frac{3\alpha_s}{16 s M_{Pl}^2} F_3\left(\frac{t}{s}, \frac{m^2}{s}\right), \quad (49.96)$$

where

$$F_1(x, y) = \frac{1}{x(y-1-x)} \left[ -4x(1+x)(1+2x+2x^2) + y(1+6x+18x^2+16x^3) - 6y^2x(1+2x) + y^3(1+4x) \right] \quad (49.97)$$

$$F_2(x, y) = -(y-1-x)F_1\left(\frac{x}{y-1-x}, \frac{y}{y-1-x}\right) \quad (49.98)$$

and

$$F_3(x, y) = \frac{1}{x(y-1-x)} \left[ 1+2x+3x^2+2x^3+x^4 - 2y(1+x^3) + 3y^2(1+x^2) - 2y^3(1+x) + y^4 \right]. \quad (49.99)$$

These formulae must then be multiplied by the graviton density of states formula  $dN = S_{n-1} \frac{M_{Pl}^2}{M_*^{n+2}} m^{n-1} dm$  to gain the cross section

$$\frac{d^2\sigma}{dtdm} = S_{n-1} \frac{M_{Pl}^2}{M_*^{n+2}} m^{n-1} \frac{d\sigma_m}{dt} \quad (49.100)$$

where  $S_n = \frac{(2\pi)^{n/2}}{\Gamma(n/2)}$  is the surface area of an  $n$ -dimensional sphere of unit radius.

Virtual graviton processes can also be searched for at colliders. For instance, in Ref. [24] the cross section for Drell-Yan production of lepton pairs via gluon fusion was calculated, where it is found that, in the center-of-mass system

$$\frac{d\sigma}{dz}(gg \rightarrow \ell^+ \ell^-) = \frac{\lambda^2 s^3}{64\pi M_*^8} (1-z^2)(1+z^2) \quad (49.101)$$

where  $z = \cos\theta$  and  $\lambda$  is a model-dependent coupling constant  $\sim 1$ . Formulae for Drell-Yan production via  $q\bar{q}$  fusion can also be found in Refs. [24,25].

#### 49.11. Warped extra dimensions

In the Randall-Sundrum model [26] of warped extra dimensions, the arena for physics is a 5-d anti-deSitter ( $AdS_5$ ) spacetime, for which a non-factorizable metric exists with a metric warp factor  $e^{-2\sigma(\phi)}$ . It is assumed that two opposite tension 3-branes exist within  $AdS_5$  at the two ends of an  $S_1/Z_2$  orbifold parametrized by co-ordinate  $\phi$  which runs from  $0 - \pi$ . The 4-D solution of the Einstein equations yields  $\sigma(\phi) = kr_c|\phi|$ , where  $r_c$  is the compactification radius of the extra dimension and  $k \sim M_{Pl}$ . The 4-D effective action allows one to identify  $\bar{M}_{Pl}^2 = \frac{M^3}{k}(1 - e^{-2kr_c\pi})$ , where  $M$  is the 5-D Planck scale. Physical particles on the TeV scale (SM) brane have mass  $m = e^{-kr_c\pi} m_0$ , where  $m_0$  is a fundamental mass of order the Planck scale. Thus, the weak scale-Planck scale hierarchy occurs due to the existence of the exponential warp factor if  $kr_c \sim 12$ .

In the simplest versions of the RS model, the TeV-scale brane contains only SM particles plus a tower of KK gravitons. The RS gravitons have mass  $m_n = kx_n e^{-kr_c\pi}$ , where the  $x_i$  are roots of Bessel functions  $J_1(x_n) = 0$ , with  $x_1 \simeq 3.83$ ,  $x_2 \simeq 7.02$  etc. While the RS zero-mode graviton couplings suppressed by  $1/\bar{M}_{Pl}$  and are thus inconsequential for collider searches, the  $n=1$  and higher modes have couplings suppressed instead by  $\Lambda_\pi = e^{-kr_c\pi} \bar{M}_{Pl} \sim TeV$ . The  $n=1$  RS graviton should have width  $\Gamma_1 = \rho m_1 x_1^2 (k/\bar{M}_{Pl})^2$ , where  $\rho$  is a constant depending on how many decay modes are open. The formulae for dilepton production via virtual RS graviton exchange can be gained from the above formulae for the ADD scenario via the replacement [27]

$$\frac{\lambda}{M_*^4} \rightarrow \frac{i^2}{8\Lambda_\pi^2} \sum_{n=1}^{\infty} \frac{1}{s - m_n^2 + im_n\Gamma_n}. \quad (49.102)$$

#### References:

1. J.F. Owens *et al.*, Phys. Rev. **D18**, 1501 (1978). Note that cross section given in previous editions of RPP for  $gg \rightarrow q\bar{q}$  lacked a factor of  $\pi$ .
2. F. Wilczek, Phys. Rev. Lett. **39**, 1304 (1977).
3. B.L. Ioffe and V.Khoze, Leningrad Report 274, 1976; Sov. J. Nucl. Phys. **9**, 50 (1978).
4. J. Ellis *et al.*, Nucl. Phys. **B106**, 292 (1976).
5. R.N. Cahn and S. Dawson, Phys. Lett. **B136**, 196 (1984), erratum, Phys. Lett. **B138**, 464 (1984).
6. S. Dawson, Nucl. Phys. **B249**, 42 (1985).
7. M.S. Chanowitz and M.K. Gaillard, Phys. Lett. **B142**, 85 (1984).
8. R.N. Cahn, Nucl. Phys. **B255**, 341 (1985).
9. For an exhaustive treatment, see V.M. Budnev *et al.*, Phys. Reports **15C**, 181(1975).
10. See *e.g.* H. Haber, *Supersymmetry, Part I (Theory)*, this review.
11. P. R. Harrison and C. H. Llewellyn Smith, Nucl. Phys. **B213**, 223 (1983), Erratum-*ibid.*, **B223**, 542 (1983); S. Dawson, E. Eichten, and C. Quigg, Phys. Rev. **D31**, 1581 (1985); V. Barger *et al.*, Phys. Rev. **D31**, 528 (1985); H. Baer and X. Tata, Phys. Lett. **B160**, 159 (1985).
12. H. Baer, D. Karatas, and X. Tata, Phys. Rev. **D42**, 2259 (1990).
13. H. Haber and G. Kane, Phys. Rept. **117**, 75 (1985).
14. Theory and Phenomenology of Sparticles, M. Drees, R. Godbole, and P. Roy (World Scientific) 2005.

15. Weak Scale Supersymmetry: From Superfields to Scattering Events, H. Baer and X. Tata (Cambridge University Press) 2006.
16. A. Bartl, H. Fraas, and W. Majerotto, *Z. Phys.* **C30**, 441 (1986).
17. H. Baer *et al.*, *Int. J. Mod. Phys.* **A4**, 4111 (1989).
18. A. Bartl, H. Fraas, and W. Majerotto, *Nucl. Phys.* **B278**, 1 (1986).
19. T. Appelquist, H.C. Cheng, and B. Dobrescu, *Phys. Rev.* **D64**, 035002 (2001).
20. For a review of models with extra spacetime dimensions, see G. Giudice and J. Wells, *Extra Dimensions*, this *Review*.
21. J.M. Smillie and B.R. Webber, *JHEP* **0510**, 069 (2005).
22. C. Macesanu, C.D. McMullen, and S. Nandi, *Phys. Rev.* **D66**, 015009 (2002).
23. N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, *Phys. Lett.* **B429**, 263 (1998) and *Phys. Rev.* **D59**, 086004 (1999).
24. J. L. Hewett, *Phys. Rev. Lett.* **82**, 4765 (1999).
25. G. Giudice, R. Rattazzi, and J. Wells, *Nucl. Phys.* **B544**, 3 (1999); E.A. Mirabeli, M. Perelstein, and M.E. Peskin, *Phys. Rev. Lett.* **82**, 2236 (1999); T. Han, J. Lykken, and R. Zhang, *Phys. Rev.* **D59**, 105006 (1999).
26. L. Randall and R.S. Sundrum, *Phys. Rev. Lett.* **83**, 3370 (1999).
27. H. Davoudiasl, J.L. Hewett, and T.G. Rizzo, *Phys. Rev. Lett.* **84**, 2080 (2000).

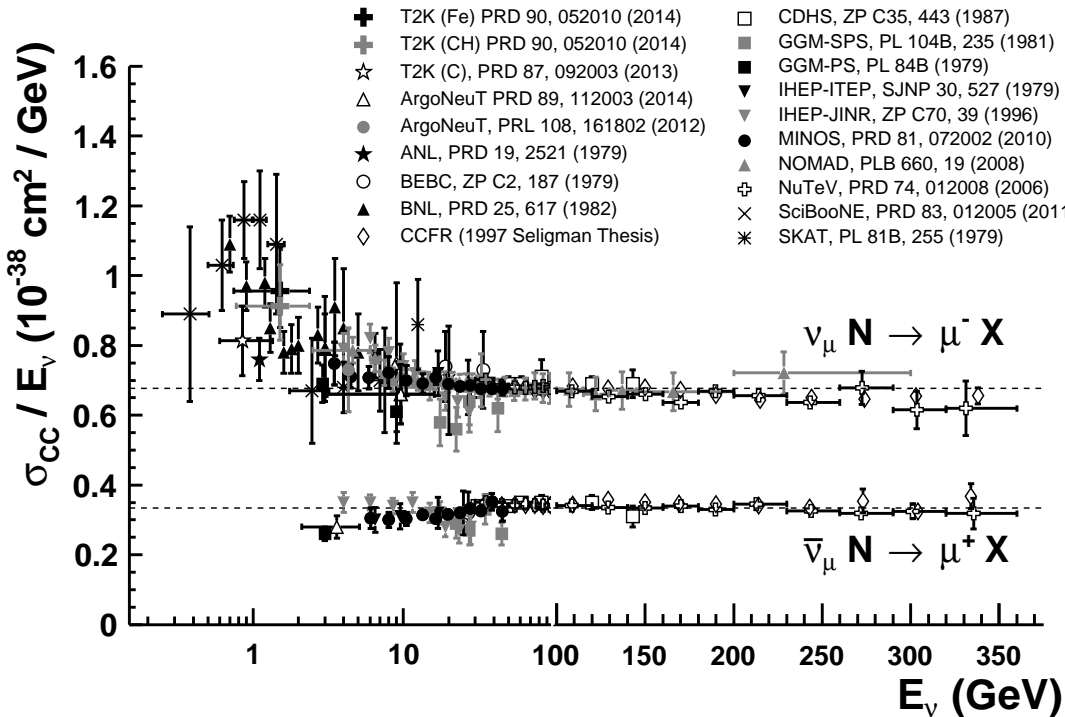
## 50. Neutrino Cross Section Measurements

Revised August 2015 by G.P. Zeller (Fermilab)

Neutrino cross sections are an essential ingredient in all neutrino experiments. Interest in neutrino scattering has recently increased due to the need for such information in the interpretation of neutrino oscillation data. Historically, neutrino scattering results on both charged current (CC) and neutral current (NC) channels have been collected over many decades using a variety of targets, analysis techniques, and detector technologies. With the advent of intense neutrino sources constructed for neutrino oscillation investigations, experiments are now remeasuring these cross sections with a renewed appreciation for nuclear effects<sup>†</sup> and the importance of improved neutrino flux estimations. This work summarizes accelerator-based neutrino cross section measurements performed in the  $\sim 0.1 - 300$  GeV range with an emphasis on inclusive, quasi-elastic, and pion production processes, areas where we have the most experimental input at present (Table 50.1 and Table 50.2). For a more comprehensive discussion of neutrino cross sections, including neutrino-electron elastic scattering and lower energy neutrino measurements, the reader is directed to a recent review of this subject [1]. Here, we survey existing experimental data on neutrino interactions and do not attempt to provide a census of the associated theoretical calculations, which are both important and plentiful.

### 50.1. Inclusive Scattering

Over the years, many experiments have measured the total inclusive cross section for neutrino ( $\nu_\mu N \rightarrow \mu^- X$ ) and antineutrino ( $\bar{\nu}_\mu N \rightarrow \mu^+ X$ ) scattering off nucleons covering a broad range of neutrino energies. As can be seen in Fig. 50.1, the inclusive cross section approaches a linear dependence on neutrino energy. Such behavior is expected for point-like scattering of neutrinos from quarks, an assumption which breaks down at lower energies. To provide a more complete picture, differential cross sections for such inclusive scattering processes have been reported – these include measurements on iron from NuTeV [42] and, more recently, at lower neutrino energies on argon from ArgoNeuT [2,3] and on carbon from T2K [37]. MINERvA has also provided new measurements of the ratios of the muon neutrino CC inclusive scattering cross section on a variety of nuclear targets such as lead, iron, and carbon [11]. At high energy, the inclusive cross section is dominated by deep inelastic scattering (DIS). Several high energy neutrino experiments have measured the DIS cross sections for specific final states, for example opposite-sign dimuon production. The most recent dimuon cross section measurements include those from CHORUS [43], NOMAD [44], and NuTeV [45]. At lower neutrino energies, the inclusive cross section is an additionally complex combination of quasi-elastic scattering and pion production processes, two areas we discuss next.



**Fig. 50.1:** Measurements of  $\nu_\mu$  and  $\bar{\nu}_\mu$  CC inclusive scattering cross sections (per nucleon) divided by neutrino energy as a function of neutrino energy. Note the transition between logarithmic and linear scales occurring at 100 GeV. Neutrino cross sections are typically twice as large as their corresponding antineutrino counterparts, although this difference can be larger at lower energies. NC cross sections (not shown) are generally smaller but non-negligible compared to the CC scattering case.

<sup>†</sup> Nuclear effects refer to kinematic and final state effects which impact neutrino scattering off nuclei. Such effects can be significant and are particularly relevant given that modern neutrino experiments make use of nuclear targets to increase their event yields.

**Table 50.1:** List of beam properties, nuclear targets, and durations for modern accelerator-based neutrino experiments studying neutrino scattering.

Experiment	beam	$\langle E_\nu \rangle, \langle E_{\bar{\nu}} \rangle$ GeV	neutrino target(s)	run period
ArgoNeuT	$\nu, \bar{\nu}$	4.3, 3.6	Ar	2009 – 2010
ICARUS	$\nu$	20.0	Ar	2010 – 2012
K2K	$\nu$	1.3	CH, H <sub>2</sub> O	2003 – 2004
MicroBooNE	$\nu$	0.8	Ar	2015 –
MINERvA	$\nu, \bar{\nu}$	3.5 (LE), 5.5 (ME)	He, CH, H <sub>2</sub> O, Fe, Pb	2009 –
MiniBooNE	$\nu, \bar{\nu}$	0.8, 0.7	CH <sub>2</sub>	2002 – 2012
MINOS	$\nu, \bar{\nu}$	3.5, 6.1	Fe	2004 –
NOMAD	$\nu, \bar{\nu}$	23.4, 19.7	C	1995 – 1998
NOvA	$\nu, \bar{\nu}$	2.0, 2.0	CH <sub>2</sub>	2010 –
SciBooNE	$\nu, \bar{\nu}$	0.8, 0.7	CH	2007 – 2008
T2K	$\nu, \bar{\nu}$	0.6, 0.6	CH, H <sub>2</sub> O	2010 –

**Table 50.2:** Summary of published neutrino cross section measurements from modern accelerator-based experiments. All measurements are  $\nu_\mu$  or  $\bar{\nu}_\mu$  scattering with the exception of the last column which is a  $\nu_e$  measurement.

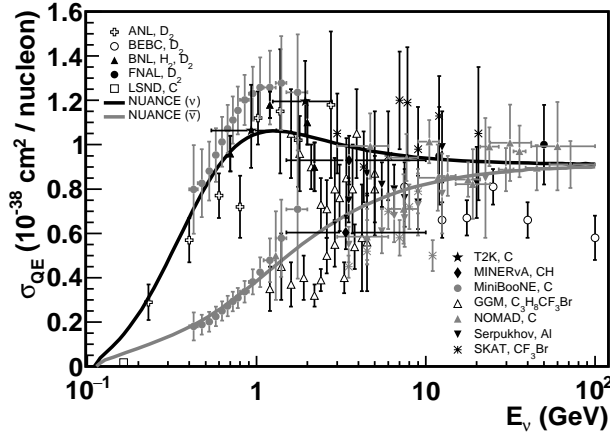
Experiment	inclusive	$0\pi$	$\pi^\pm$	$\pi^0$	$\nu_e$
ArgoNeuT	CC [2,3]	2p [4]	CC [5]	–	–
K2K	–	CC [6]	CC [7,8]	CC [9], NC [10]	–
MINERvA	CC [11]	CC [12,13], 1p [14]	CC [15]	CC [16]	–
MiniBooNE	–	CC [17,18], $M_A$ [19], NC [20,21,22]	CC [23,24]	CC [25], NC [26,27]	–
MINOS	CC [28]	$M_A$ [29]	–	–	–
NOMAD	CC [30]	CC [31]	–	NC [32]	–
SciBooNE	CC [33]	–	CC [34]	NC [35,36]	–
T2K	CC [37,38]	CC [39], NC [40]	–	–	CC [41]

## 50.2. Quasi-elastic scattering

Quasi-elastic (QE) scattering is the dominant neutrino interaction for neutrino energies less than  $\sim 1$  GeV and represents a large fraction of the signal samples in many neutrino oscillation experiments. Historically, neutrino (antineutrino) quasi-elastic scattering refers to the process,  $\nu_\mu n \rightarrow \mu^- p$  ( $\bar{\nu}_\mu p \rightarrow \mu^+ n$ ), where a charged lepton and single nucleon are ejected in the elastic interaction of a neutrino (or antineutrino) with a nucleon in the target material. This is the final state one would strictly observe, for example, in scattering off of a free nucleon target. Fig. 50.2 displays the current status of existing measurements of  $\nu_\mu$  and  $\bar{\nu}_\mu$  QE scattering cross sections as a function of neutrino energy. In this plot, and all others in this review, the prediction from a representative neutrino event generator (NUANCE) [46] provides a theoretical comparator. Other generators and more sophisticated calculations exist which can yield significantly different predictions [47]. Note that modern experiments have recently opted to report QE cross sections as a function of final state muon or proton kinematics [17,18,48]. Such distributions are more difficult to compare between experiments but are much less model-dependent and provide more stringent tests of the theory than cross sections as a function of neutrino energy ( $E_\nu$ ) or 4-momentum transfer ( $Q^2$ ).

In many of these initial measurements of the neutrino QE cross section, bubble chamber experiments employed light targets ( $H_2$  or  $D_2$ ) and required both the detection of the final state muon and single nucleon<sup>‡</sup>; thus the final state was clear and elastic kinematic conditions could be verified. The situation is more complicated, of course, for heavier nuclear targets. In this case, nuclear effects can impact the size and shape of the cross section as well as the final state kinematics and topology. Due to intranuclear hadron rescattering and the possible effects of correlations between target nucleons, additional nucleons may be ejected in the final state; hence, a QE interaction on a nuclear target does not always imply the ejection of a *single* nucleon. One therefore needs to take some care in defining what one means by neutrino QE scattering when scattering off targets heavier than  $H_2$  or  $D_2$ . Adding to the complexity, recent MiniBooNE measurements of the  $\nu_\mu$  and  $\bar{\nu}_\mu$  QE scattering cross sections on carbon near 1 GeV have revealed a significantly larger cross section than originally anticipated [17,18]. Such an enhancement was observed many years prior in electron-nucleus scattering [57] and is believed to be due to the presence of correlations between target nucleons in the nucleus. As a result, the impact of such nuclear effects on neutrino QE scattering has recently

<sup>‡</sup> In the case of  $D_2$ , many experiments additionally observed the spectator proton.



**Figure 50.2:** Measurements of  $\nu_\mu$  (black) and  $\bar{\nu}_\mu$  (red) QE scattering cross sections (per nucleon) as a function of neutrino energy. Data on a variety of nuclear targets are shown, including measurements from ANL [49], BEBC [50], BNL [51], FNAL [52], LSND [53], T2K [39], MINERvA [12], MiniBooNE [17,18], GGM [54], NOMAD [31], Serpukhov [55], and SKAT [56]. Shown is the QE free nucleon scattering prediction from NUANCE [46] assuming  $M_A = 1.0$  GeV. This prediction is significantly altered by nuclear effects in the case of neutrino-nucleus scattering. Although plotted together, care should be taken in interpreting measurements performed on targets heavier than  $D_2$  due to possible differences in QE identification and kinematics.

been the subject of intense experimental and theoretical scrutiny with potential implications on event rates, nucleon emission, neutrino energy reconstruction, and neutrino/antineutrino ratios. The reader is referred to a recent review of the situation in [58,59]. Additional measurements are clearly needed before a complete understanding is achieved. To help drive further progress, neutrino-nucleus QE cross sections have been reported for the first time in the form of double-differential distributions in muon kinematics,  $d^2\sigma/dT_\mu d\cos\theta_\mu$ , by MiniBooNE [17,18] thus reducing the model-dependence of the reported data and allowing a more rigorous two-dimensional test of the underlying nuclear theory. Experiments such as ArgoNeuT have begun to provide the first measurements of proton multiplicities in neutrino-argon QE scattering [4,48], a critical ingredient in understanding the hadronic side of these interactions and final state effects. MINOS, NOvA, and T2K have started to study QE interactions in their near detectors with sizable statistics [29,39,60]. Most recently, MINERvA has measured the differential cross section,  $d\sigma/dQ_{QE}^2$ ,  $\nu_e$  QE scattering [61], single proton emission [14], and vertex energy in both  $\nu_\mu$  and  $\bar{\nu}_\mu$  QE interactions in hydrocarbon [12,13]. With the MiniBooNE results having revealed additional complexities in the QE channel, measurements from other neutrino experiments are crucial for getting a better handle on the underlying nuclear physics impacting neutrino-nucleus interactions. What we once thought was “simple” QE scattering is in fact not so simple.

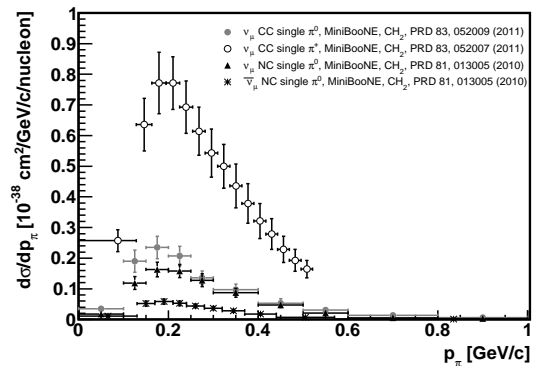
In addition to such charged current investigations, measurements of the neutral current counterpart of this channel have also been performed. The most recent NC elastic scattering cross section measurements include those from BNL E734 [62], MiniBooNE [20,21,22], and T2K [40]. A number of measurements of the Cabibbo-suppressed antineutrino QE hyperon production cross section have additionally been reported [56,63], although not in recent years.

### 50.3. Pion Production

In addition to such elastic processes, neutrinos can also inelastically scatter producing a nucleon excited state ( $\Delta$ ,  $N^*$ ). Such baryonic resonances quickly decay, most often to a nucleon and single-pion final state. Historically, experiments have measured various exclusive final states associated with these reactions, the majority of which have been on hydrogen and deuterium targets [1].

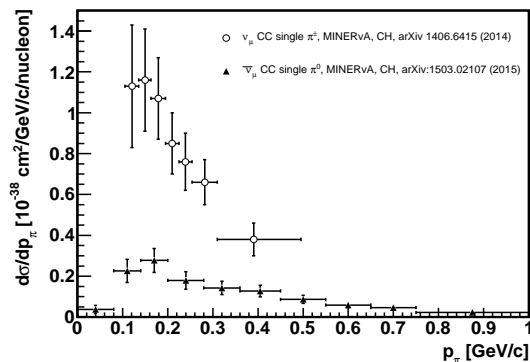
In addition to such resonance production processes, neutrinos can also coherently scatter off of the entire nucleus and produce a distinctly forward-scattered single pion final state. Both CC ( $\nu_\mu A \rightarrow \mu^- A\pi^+$ ,  $\bar{\nu}_\mu A \rightarrow \mu^+ A\pi^-$ ) and NC ( $\nu_\mu A \rightarrow \nu_\mu A\pi^0$ ,  $\bar{\nu}_\mu A \rightarrow \bar{\nu}_\mu A\pi^0$ ) processes are possible in this case. Even though the level of coherent pion production is small compared to resonant processes, observations exist across a broad energy range and on multiple nuclear targets [64]. More recently, several modern neutrino experiments have measured coherent pion production cross sections including ArgoNeuT [5], K2K [8], MINERvA [15], MiniBooNE [27], MINOS [65], NOMAD [32], and SciBooNE [34,36].

As with QE scattering, a new appreciation for the significance of nuclear effects has surfaced in pion production channels, again due to the use of heavy nuclear targets in modern neutrino experiments. Many experiments have been careful to report cross sections for various detected final states, thereby not correcting for large and uncertain nuclear effects (e.g., pion rescattering, charge exchange, and absorption) which can introduce unwanted sources of uncertainty and model dependence. Recent measurements of single-pion cross sections, as published by K2K [9,10], MiniBooNE [24], and SciBooNE [35], take the form of ratios with respect to QE or CC inclusive scattering samples. Providing the most comprehensive survey of neutrino single-pion production to date, MiniBooNE has recently published a total of 16 single- and double-differential cross sections for both the final state muon (in the case of CC scattering) and pion in these interactions; thus, providing the first measurements of these distributions (Fig. 50.3) [23–26]. MINERvA has recently produced similar kinematic measurements at higher neutrino energies (Fig. 50.4) [16,66]. Regardless of the interaction channel, such differential cross section measurements (in terms of observed final state particle kinematics) are now preferred for their reduced model dependence and for the additional kinematic information they provide. Such a new direction has been the focus of modern measurements as opposed to the reporting of more model-dependent, historical cross sections as a function of  $E_\nu$  or  $Q^2$ . Together with similar results for other interaction channels, a better understanding and modeling of nuclear effects will be possible moving forward.



**Figure 50.3:** Differential cross sections for CC and NC pion production from MiniBooNE at a mean neutrino energy of 0.8 GeV. Shown here are the measurements as a function of the momentum of the outgoing pion in the interaction, a kinematic that is particularly sensitive to final state interactions. Other distributions are also available in the publications listed in the legend.

It should be noted that baryonic resonances can also decay to multi-pion, other mesonic ( $K$ ,  $\eta$ ,  $\rho$ , etc.), and even photon final states. Experimental results for these channels are typically sparse or non-existent [1]; however, photon production processes can be an important background for  $\nu_\mu \rightarrow \nu_e$  appearance searches and thus have become the focus of some recent experimental investigations; for example, in NOMAD [67].



**Figure 50.4:** Differential cross sections for neutrino and antineutrino pion production from MINERvA at a mean neutrino energy of 3.3 GeV. Shown here are the measurements as a function of the momentum of the outgoing pion in the interaction, a kinematic that is particularly sensitive to final state interactions. Other distributions are also available in the publications listed in the legend. Note that while the MINERvA  $\nu_\mu$  measurement includes both  $\pi^+$  and  $\pi^-$  production, the sample is almost entirely ( $> 99\%$ )  $\pi^+$  final states.

#### 50.4. Outlook

Currently operating experiments will continue to produce additional neutrino cross section measurements as they accumulate additional statistics, while a few new experiments will soon be coming online. In the coming years, analysis of a broad energy range of data on a variety of targets in the MINERvA experiment will provide the most detailed analysis yet of nuclear effects in neutrino interactions. Data from ArgoNeuT, ICARUS, MicroBooNE, and SBND will probe deeper into complex neutrino final states using the superior capabilities of liquid argon time projection chambers, while the T2K and NOvA near detectors will collect high statistics samples in intense neutrino beams. Together, these investigations should significantly advance our understanding of neutrino-nucleus scattering in the next decade.

#### 50.5. Acknowledgments

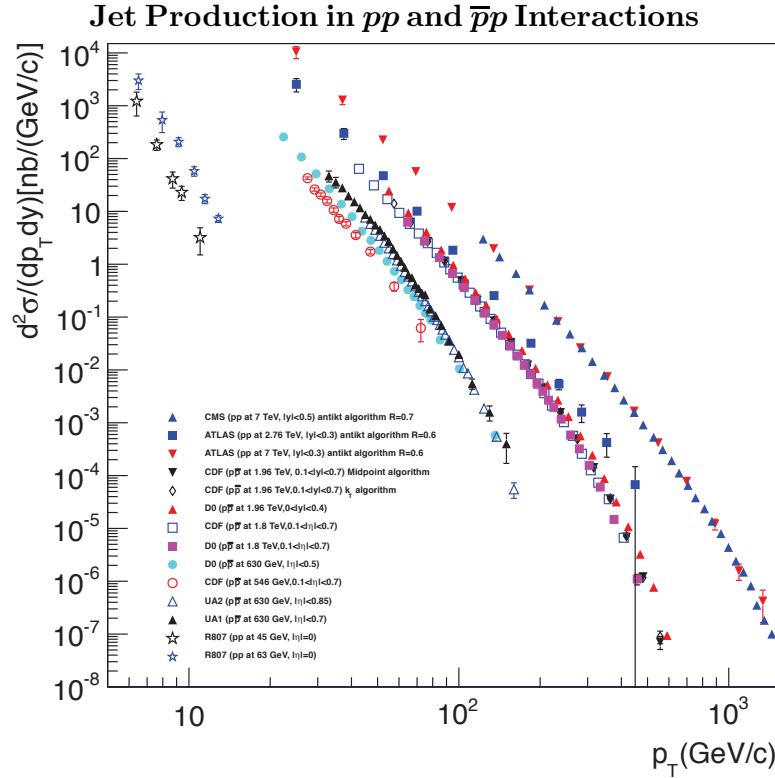
The author thanks Anne Schukraft (Fermilab) for help in updating the plots contained in this review.

#### References:

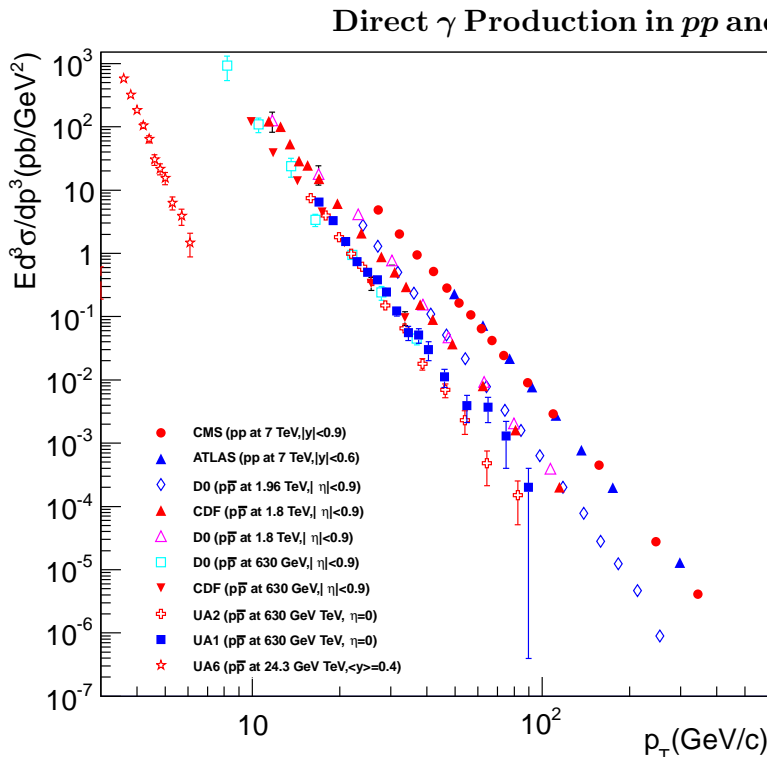
- J.A. Formaggio and G.P. Zeller, *Rev. Mod. Phys.* **84**, 1307 (2012).
- C. Anderson *et al.*, *Phys. Rev. Lett.* **108**, 161802 (2012).
- R. Acciarri *et al.*, *Phys. Rev.* **D89**, 112003 (2014).
- R. Acciarri *et al.*, *Phys. Rev.* **D90**, 012008 (2014).
- R. Acciarri *et al.*, *Phys. Rev. Lett.* **113**, 261801 (2014).
- R. Gran *et al.*, *Phys. Rev.* **D74**, 052002 (2006).
- A. Rodriguez *et al.*, *Phys. Rev.* **D78**, 032003 (2008).
- M. Hasegawa *et al.*, *Phys. Rev. Lett.* **95**, 252301 (2005).
- C. Mariani *et al.*, *Phys. Rev.* **D83**, 054023 (2011).
- S. Nakayama *et al.*, *Phys. Lett.* **B619**, 255 (2005).
- B.G. Tice *et al.*, *Phys. Rev. Lett.* **112**, 231801 (2014).
- G.A. Fiorentini *et al.*, *Phys. Rev. Lett.* **111**, 022502 (2013).
- L. Fields *et al.*, *Phys. Rev. Lett.* **111**, 022501 (2013).
- T. Walton *et al.*, *Phys. Rev.* **D91**, 071301 (2015).
- A. Higuera *et al.*, *Phys. Rev. Lett.* **113**, 261802 (2014).
- T. Le *et al.*, *Phys. Lett.* **B749**, 130 (2015).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D81**, 092005 (2010).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D88**, 032001 (2013).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev. Lett.* **100**, 032301 (2008).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D82**, 092005 (2010).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D91**, 012004 (2015).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D91**, 012004 (2015).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D83**, 052007 (2011).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev. Lett.* **103**, 081801 (2009).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D83**, 052009 (2011).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Rev.* **D81**, 013005 (2010).
- A.A. Aguilar-Arevalo *et al.*, *Phys. Lett.* **B664**, 41 (2008).
- P. Adamson *et al.*, *Phys. Rev.* **D81**, 072002 (2010).
- P. Adamson *et al.*, *Phys. Rev.* **91**, 012005 (2015).
- Q. Wu *et al.*, *Phys. Lett.* **B660**, 19 (2008).
- V. Lyubushkin *et al.*, *Eur. Phys. J.* **C63**, 355 (2009).
- C.T. Kullenberg *et al.*, *Phys. Lett.* **B682**, 177 (2009).
- Y. Nakajima *et al.*, *Phys. Rev.* **D83**, 12005 (2011).
- K. Hiraide *et al.*, *Phys. Rev.* **D78**, 112004 (2008).
- Y. Kurimoto *et al.*, *Phys. Rev.* **D81**, 033004 (2010).
- Y. Kurimoto *et al.*, *Phys. Rev.* **D81**, 111102 (R)(2010).
- K. Abe *et al.*, *Phys. Rev.* **D87**, 092003 (2013).
- K. Abe *et al.*, *Phys. Rev.* **D90**, 052010 (2014).
- K. Abe *et al.*, *Phys. Rev.* **D91**, 112002 (2015).
- K. Abe *et al.*, *Phys. Rev.* **D90**, 072012 (2014).
- K. Abe *et al.*, *Phys. Rev. Lett.* **113**, 241803 (2014).
- M. Tzanov *et al.*, *Phys. Rev.* **D74**, 012008 (2006).
- A. Kayis-Topaksu *et al.*, *Nucl. Phys.* **B798**, 1 (2008).
- O. Samoylov *et al.*, *Nucl. Phys.* **B876**, 339 (2013).
- D. Mason *et al.*, *Phys. Rev. Lett.* **99**, 192001 (2007).
- D. Casper, *Nucl. Phys. (Proc. Supp.)* **112**, 161 (2002), default v3 NUANCE.
- R. Tacik, *AIP Conf. Proc.* **1405**, 229 (2011); S. Boyd *et al.*, *AIP Conf. Proc.* **1189**, 60 (2009).
- O. Palamara *et al.*, *arXiv:1309.7480* [physics.ins-det].
- S.J. Barish *et al.*, *Phys. Rev.* **D16**, 3103 (1977).
- D. Allasia *et al.*, *Nucl. Phys.* **B343**, 285 (1990).
- N.J. Baker *et al.*, *Phys. Rev.* **D23**, 2499 (1981); G. Fanourakis *et al.*, *Phys. Rev.* **D21**, 562 (1980).
- T. Kitagaki *et al.*, *Phys. Rev.* **D28**, 436 (1983).
- L.B. Auerbach *et al.*, *Phys. Rev.* **C66**, 015501 (2002).
- S. Bonetti *et al.*, *Nuovo Cimento* **A38**, 260 (1977); N. Armenise *et al.*, *Nucl. Phys.* **B152**, 365 (1979).
- S.V. Belikov *et al.*, *Z. Phys.* **A320**, 625 (1985).
- J. Brunner *et al.*, *Z. Phys.* **C45**, 551 (1990).
- J. Carlson *et al.*, *Phys. Rev.* **C65**, 024002 (2002).
- H. Gallagher *et al.*, *Ann. Rev. Nucl. and Part. Sci.* **61**, 355 (2011).
- G.T. Garvey *et al.*, *Phys. Reports* **580**, 1 (2015).
- M. Betancourt, Ph.D. thesis, University of Minnesota, 2013.
- J. Wolcott *et al.*, *arXiv:1509.05729* [hep-ex].
- L.A. Ahrens *et al.*, *Phys. Rev.* **D35**, 785 (1987).
- V.V. Ammosov *et al.*, *Z. Phys.* **C36**, 377 (1987); O. Erriques *et al.*, *Phys. Lett.* **70B**, 383 (1977); T. Eichten *et al.*, *Phys. Lett.* **40B**, 593 (1972).
- For a compilation of historical coherent pion production data, please see P. Villain *et al.*, *Phys. Lett.* **B313**, 267 (1993).
- D. Cherdack, *AIP Conf. Proc.* **1405**, 115 (2011).
- B. Eberly *et al.*, *Phys. Rev.* **D92**, 092008 (2015).
- C.T. Kullenberg *et al.*, *Phys. Lett.* **B706**, 268 (2012).

## 51. PLOTS OF CROSS SECTIONS AND RELATED QUANTITIES

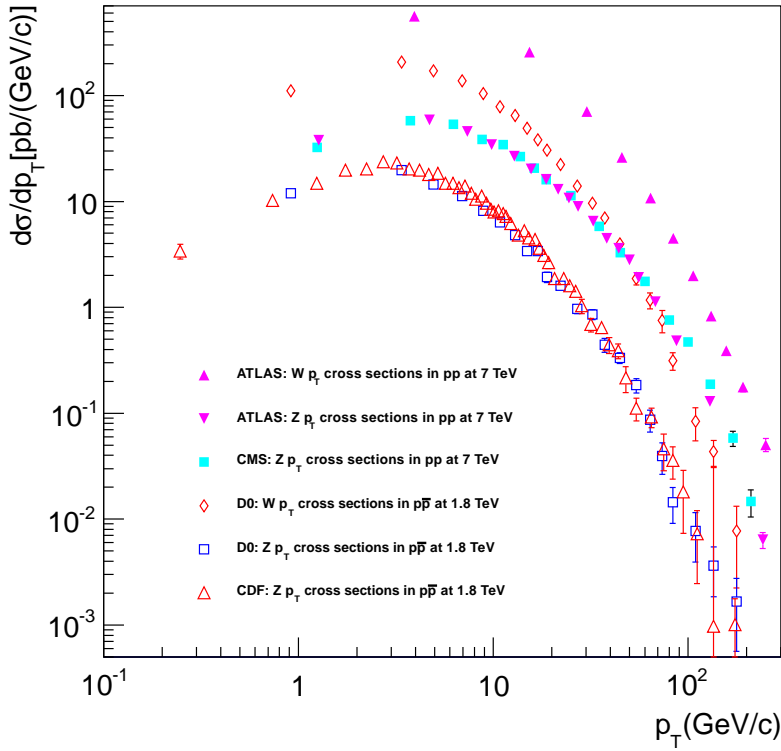
(For neutrino plots, see review article "Neutrino Cross Section Measurements" by G.P. Zeller in this edition of RPP)



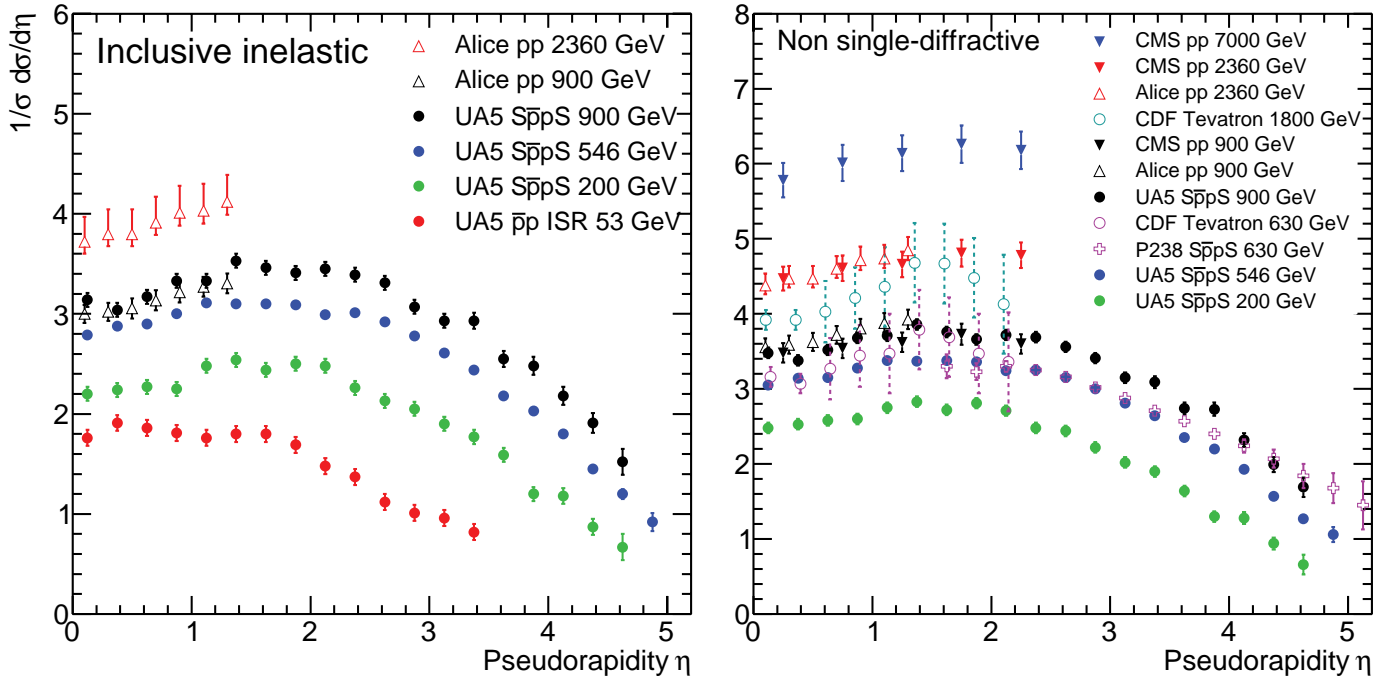
**Figure 51.1:** Inclusive differential jet cross sections, in the central rapidity region, plotted as a function of the jet transverse momentum. Results earlier than from the Tevatron Run 2 used transverse energy rather than transverse momentum and pseudo-rapidity  $\eta$  rather than rapidity  $y$ , but  $p_T$  and  $y$  are used for all results shown here for simplicity. The error bars plotted are in most cases the experimental stat. and syst. errors added in quadrature. The CDF and D0 measurements use jet sizes of 0.7 (JetClu for CDF Run 1, and Midpoint and  $k_T$  for CDF Run 2, a cone algorithm for D0 in Run 1 and the Midpoint algorithm in Run 2). The ATLAS results are plotted for the antikt algorithm for  $R=0.4$ , while the CMS results also use antikt, but with  $R=0.5$ . NLO QCD predictions in general provide a good description of the Tevatron and LHC data; the Tevatron jet data in fact are crucial components of global PDF fits, and the LHC data are starting to be used as well. Comparisons with the older cross sections are more difficult due to the nature of the jet algorithms used. **ATLAS:** Phys. Rev. **D86**, 014022 (2012), Eur. Phys. J **C73**, 2509 (2013); **CMS:** Phys. Rev. **D84**, 052011 (2011); **CDF:** Phys. Rev. **D75**, 092006 (2007), Phys. Rev. **D64**, 032001 (2001), Phys. Rev. Lett. **70**, 1376 (1993); **D0:** Phys. Rev. **D64**, 032003 (2001); **UA2:** Phys. Lett. **B257**, 232 (1991); **UA1:** Phys. Lett. **172**, 461 (1986); **R807:** Phys. Lett. **B123**, 133 (1983). (Courtesy of J. Huston, Michigan State University, 2013.)



**Figure 51.2:** Isolated photon cross sections plotted as a function of the photon transverse momentum. The errors are either statistical only, or statistical and systematic added in quadrature. **ATLAS:** Phys. Lett. **B706**, 150 (2011); **CMS:** Phys. Rev. **D84**, 052011 (2011); **D0:** Phys. Lett. **B639**, 151 (2006), Phys. Rev. Lett. **87**, 251805 (2001); **CDF:** Phys. Rev. **D65**, 112003 (2002); **UA6:** Phys. Lett. **B206**, 163 (1988); **UA1:** Phys. Lett. **B209**, 385 (1988); **UA2:** Phys. Lett. **B288**, 386 (1992). (Courtesy of J. Huston, Michigan State University, 2013.)

Differential Cross Section for  $W$  and  $Z$  Boson Production

**Figure 51.3:** Differential cross sections for  $W$  and  $Z$  production shown as a function of the boson transverse momentum. The errors plotted are either statistical only or statistical and systematic added in quadrature. The results are in good agreement with theoretical predictions that include both the effects of NLO corrections and of  $q_T$  resummation. **ATLAS:** Phys. Rev. **D85**, 012005 (2012), Phys. Lett. **B705**, 415 (2011); **CMS:** Phys. Rev. **D85**, 032002 (2012); **D0:** Phys. Lett. **B513**, 292 (2001), Phys. Rev. Lett. **84**, 2792 (2000); **CDF:** Phys. Rev. Lett. **84**, 845 (2000). (Courtesy of J. Huston, Michigan State University, 2013.)

Pseudorapidity Distributions in  $pp$  and  $p\bar{p}$  Interactions

**Figure 51.4:** Charged particle pseudorapidity distributions in  $pp$  collisions for  $53 \text{ GeV} \leq \sqrt{s} \leq 1800 \text{ GeV}$ . UA5 data from the  $S\bar{p}p$ S are taken from G.J. Alner *et al.*, Z. Phys. **C33**, 1 (1986), and from the ISR from K. Alpgöard *et al.*, Phys. Lett. **112B** 193 (1982). The UA5 data are shown for both the full inelastic cross-section and with singly diffractive events excluded. Additional non single-diffractive measurements are available from CDF at the Tevatron, F. Abe *et al.*, Phys. Rev. **D41**, 2330 (1990) and from P238 at the  $S\bar{p}p$ S, R. Harr *et al.*, Phys. Lett. **B401**, 176 (1997). These may be compared with both inclusive and non single-diffractive measurements in  $pp$  collisions at the LHC from ALICE, K. Aamodt *et al.*, Eur. Phys. J. **C68**, 89 (2010) and for non single-diffractive interactions from CMS, V. Khachatryan *et al.*, JHEP 1002:041 (2010), Phys. Rev. Lett. **105**, 022002 (2010). (Courtesy of D.R. Ward, Cambridge Univ., 2013)



### Average Hadron Multiplicities in Hadronic $e^+e^-$ Annihilation Events

**Table 51.1:** Average hadron multiplicities per hadronic  $e^+e^-$  annihilation event at  $\sqrt{s} \approx 10, 29\text{--}35, 91$ , and  $130\text{--}200$  GeV. The rates given include decay products from resonances with  $c\tau < 10$  cm, and include the corresponding anti-particle state. Correlations of the systematic uncertainties were considered for the calculation of the averages. Quoted errors are not increased by scale factor  $S$ . (Updated August 2015 by O. Biebel, LMU, Munich)

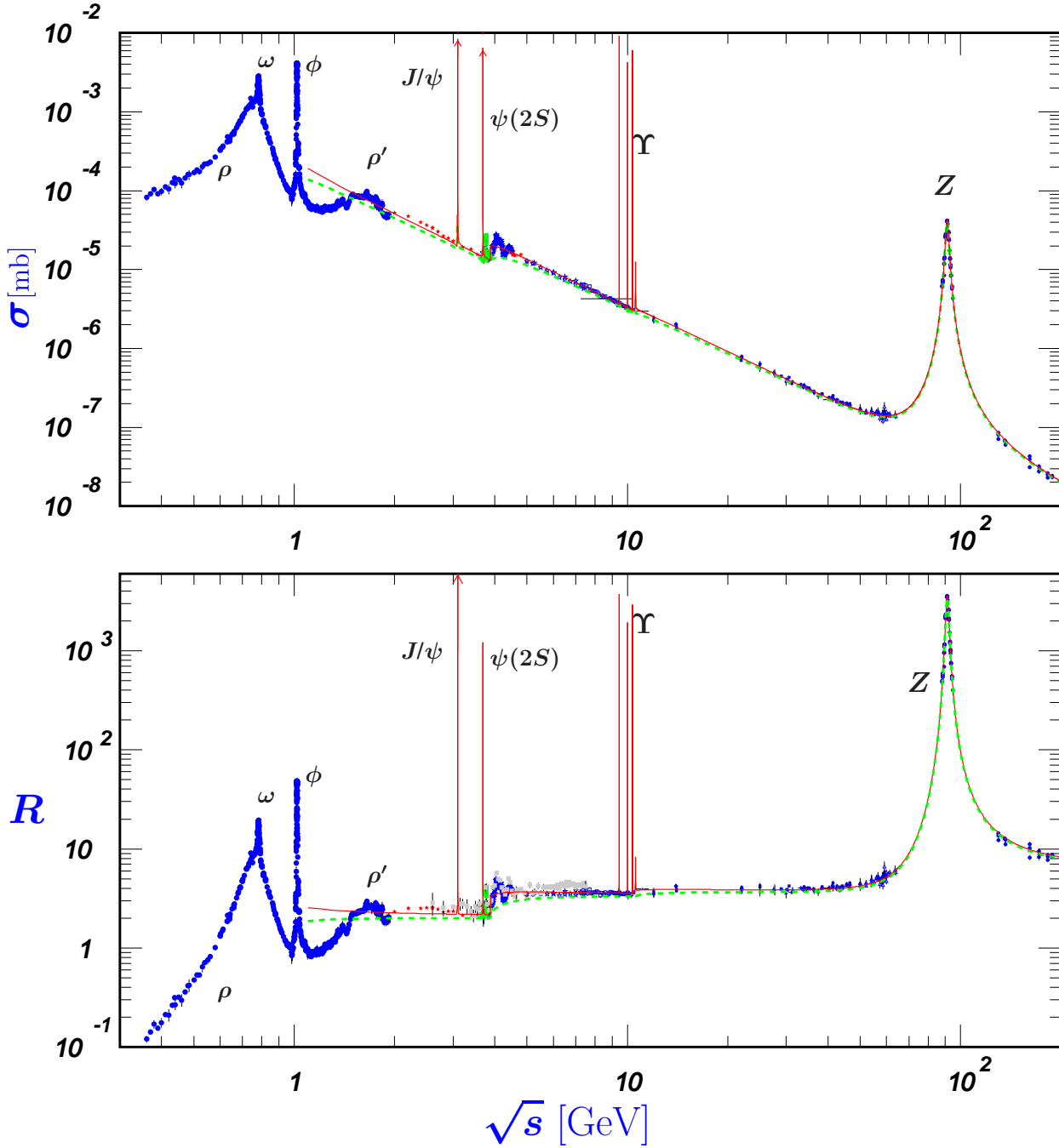
Particle	$\sqrt{s} \approx 10$ GeV	$\sqrt{s} = 29\text{--}35$ GeV	$\sqrt{s} = 91$ GeV	$\sqrt{s} = 130\text{--}200$ GeV
<b>Pseudoscalar mesons:</b>				
$\pi^+$	$6.52 \pm 0.11$	$10.3 \pm 0.4$	$17.02 \pm 0.19$	$21.24 \pm 0.39$
$\pi^0$	$3.2 \pm 0.3$	$5.83 \pm 0.28$	$9.42 \pm 0.32$	
$K^+$	$0.953 \pm 0.018$	$1.48 \pm 0.09$	$2.228 \pm 0.059$	$2.82 \pm 0.19$
$K^0$	$0.91 \pm 0.05$	$1.48 \pm 0.07$	$2.049 \pm 0.026$	$2.10 \pm 0.12$
$\eta$	$0.20 \pm 0.04$	$0.61 \pm 0.07$	$1.049 \pm 0.080$	
$\eta(958)$	$0.03 \pm 0.01$	$0.26 \pm 0.10$	$0.152 \pm 0.020$	
$D^+$	$0.194 \pm 0.019^{(a)}$	$0.17 \pm 0.03$	$0.175 \pm 0.016$	
$D^0$	$0.446 \pm 0.032^{(a)}$	$0.45 \pm 0.07$	$0.454 \pm 0.030$	
$D_s^+$	$0.063 \pm 0.014^{(a)}$	$0.45 \pm 0.20^{(b)}$	$0.131 \pm 0.021$	
$B^{(c)}$	—	—	$0.134 \pm 0.016^{(d)}$	
$B^+$	—	—	$0.141 \pm 0.004^{(d)}$	
$B_s^0$	—	—	$0.054 \pm 0.011^{(d)}$	
<b>Scalar mesons:</b>				
$f_0(980)$	$0.024 \pm 0.006$	$0.05 \pm 0.02^{(e)}$	$0.146 \pm 0.012$	
$a_0(980)^\pm$	—	—	$0.27 \pm 0.11^{(f)}$	
<b>Vector mesons:</b>				
$\rho(770)^0$	$0.35 \pm 0.04$	$0.81 \pm 0.08$	$1.231 \pm 0.098$	
$\rho(770)^\pm$	—	—	$2.40 \pm 0.43^{(f)}$	
$\omega(782)$	$0.30 \pm 0.08$	—	$1.016 \pm 0.065$	
$K^*(892)^+$	$0.27 \pm 0.03$	$0.64 \pm 0.05$	$0.714 \pm 0.055$	
$K^*(892)^0$	$0.29 \pm 0.03$	$0.56 \pm 0.06$	$0.738 \pm 0.024$	
$\phi(1020)$	$0.044 \pm 0.003$	$0.085 \pm 0.011$	$0.0963 \pm 0.0032$	
$D^*(2010)^+$	$0.177 \pm 0.022^{(a)}$	$0.43 \pm 0.07$	$0.1937 \pm 0.0057^{(g)}$	
$D^*(2007)^0$	$0.168 \pm 0.019^{(a)}$	$0.27 \pm 0.11$	—	
$D_s^*(2112)^+$	$0.048 \pm 0.014^{(a)}$	—	$0.101 \pm 0.048^{(h)}$	
$B^* \text{ }^{(i)}$	—	—	$0.288 \pm 0.026$	
$J/\psi(1S)$	$0.00050 \pm 0.00005^{(a)}$	—	$0.0052 \pm 0.0004^{(j)}$	
$\psi(2S)$	—	—	$0.0023 \pm 0.0004^{(j)}$	
$\Upsilon(1S)$	—	—	$0.00014 \pm 0.00007^{(j)}$	
<b>Pseudovector mesons:</b>				
$f_1(1285)$	—	—	$0.165 \pm 0.051$	
$f_1(1420)$	—	—	$0.056 \pm 0.012$	
$\chi_{c1}(3510)$	—	—	$0.0041 \pm 0.0011^{(j)}$	
<b>Tensor mesons:</b>				
$f_2(1270)$	$0.09 \pm 0.02$	$0.14 \pm 0.04$	$0.166 \pm 0.020$	
$f_2'(1525)$	—	—	$0.012 \pm 0.006$	
$K_2^*(1430)^+$	—	$0.09 \pm 0.03$	—	
$K_2^*(1430)^0$	—	$0.12 \pm 0.06$	$0.084 \pm 0.022$	
$B^{**} \text{ }^{(k)}$	—	—	$0.118 \pm 0.024$	
$D_{s1}^\pm$	—	—	$0.0052 \pm 0.0011^{(\ell)}$	
$D_{s2}^{*\pm}$	—	—	$0.0083 \pm 0.0031^{(\ell)}$	
<b>Baryons:</b>				
$p$	$0.266 \pm 0.008$	$0.640 \pm 0.050$	$1.050 \pm 0.032$	$1.41 \pm 0.18$
$\Lambda$	$0.080 \pm 0.007$	$0.205 \pm 0.010$	$0.3915 \pm 0.0065$	$0.39 \pm 0.03$
$\Sigma^0$	$0.023 \pm 0.008$	—	$0.078 \pm 0.010$	
$\Sigma^-$	—	—	$0.081 \pm 0.010$	
$\Sigma^+$	—	—	$0.107 \pm 0.011$	
$\Sigma^\pm$	—	—	$0.174 \pm 0.009$	
$\Xi^-$	$0.0059 \pm 0.0007$	$0.0176 \pm 0.0027$	$0.0262 \pm 0.0009$	
$\Delta(1232)^{++}$	$0.040 \pm 0.010$	—	$0.085 \pm 0.014$	
$\Sigma(1385)^-$	$0.006 \pm 0.002$	$0.017 \pm 0.004$	$0.0240 \pm 0.0017$	
$\Sigma(1385)^+$	$0.005 \pm 0.001$	$0.017 \pm 0.004$	$0.0239 \pm 0.0015$	
$\Sigma(1385)^\pm$	$0.0106 \pm 0.0020$	$0.033 \pm 0.008$	$0.0472 \pm 0.0027$	
$\Xi(1530)^0$	$0.0015 \pm 0.0006$	—	$0.00694 \pm 0.00049$	
$\Omega^-$	$0.0007 \pm 0.0004$	$0.014 \pm 0.007$	$0.00124 \pm 0.00018$	
$\Lambda_c^+$	$0.074 \pm 0.031^{(m)}$	$0.110 \pm 0.050$	$0.078 \pm 0.017$	
$\Lambda_b^0$	—	—	$0.031 \pm 0.016$	
$\Sigma_c^{++}, \Sigma_c^0$	$0.014 \pm 0.007$	—	—	
$\Lambda(1520)$	$0.008 \pm 0.002$	—	$0.0222 \pm 0.0027$	

## Notes for Table 51.1:

- (a)  $\sigma_{\text{had}} = 3.33 \pm 0.05 \pm 0.21$  nb (CLEO: Phys. Rev. **D29**, 1254 (1984)) has been used in converting the measured cross sections to average hadron multiplicities.
- (b)  $B(D_s \rightarrow \eta\pi, \eta'\pi)$  was used (RPP 1994).
- (c) Comprises both charged and neutral  $B$  meson states.
- (d) The Standard Model  $B(Z \rightarrow b\bar{b}) = 0.217$  was used.
- (e)  $x_p = p/p_{\text{beam}} > 0.1$  only.
- (f) Both charge states.
- (g)  $B(D^*(2010)^+ \rightarrow D^0\pi^+) \times B(D^0 \rightarrow K^-\pi^+)$  has been used (RPP 2000).
- (h)  $B(D_s^* \rightarrow D_s^+\gamma)$ ,  $B(D_s^+ \rightarrow \phi\pi^+)$ ,  $B(\phi \rightarrow K^+K^-)$  have been used (RPP 1998).
- (i) Any charge state (*i.e.*,  $B_d^*$ ,  $B_u^*$ , or  $B_s^*$ ).
- (j)  $B(Z \rightarrow \text{hadrons}) = 0.699$  was used (RPP 1994).
- (k) Any charge state (*i.e.*,  $B_d^{**}$ ,  $B_u^{**}$ , or  $B_s^{**}$ ).
- (l) Assumes  $B(D_{s1}^+ \rightarrow D^{*+}K^0 + D^{*0}K^+) = 100\%$  and  $B(D_{s2}^+ \rightarrow D^0K^+) = 45\%$ .
- (m) The value was derived from the cross section of  $\Lambda_c^+ \rightarrow p\pi K$  using (a) and assuming the branching fraction to be  $(5.0 \pm 1.3)\%$  (RPP 2004).

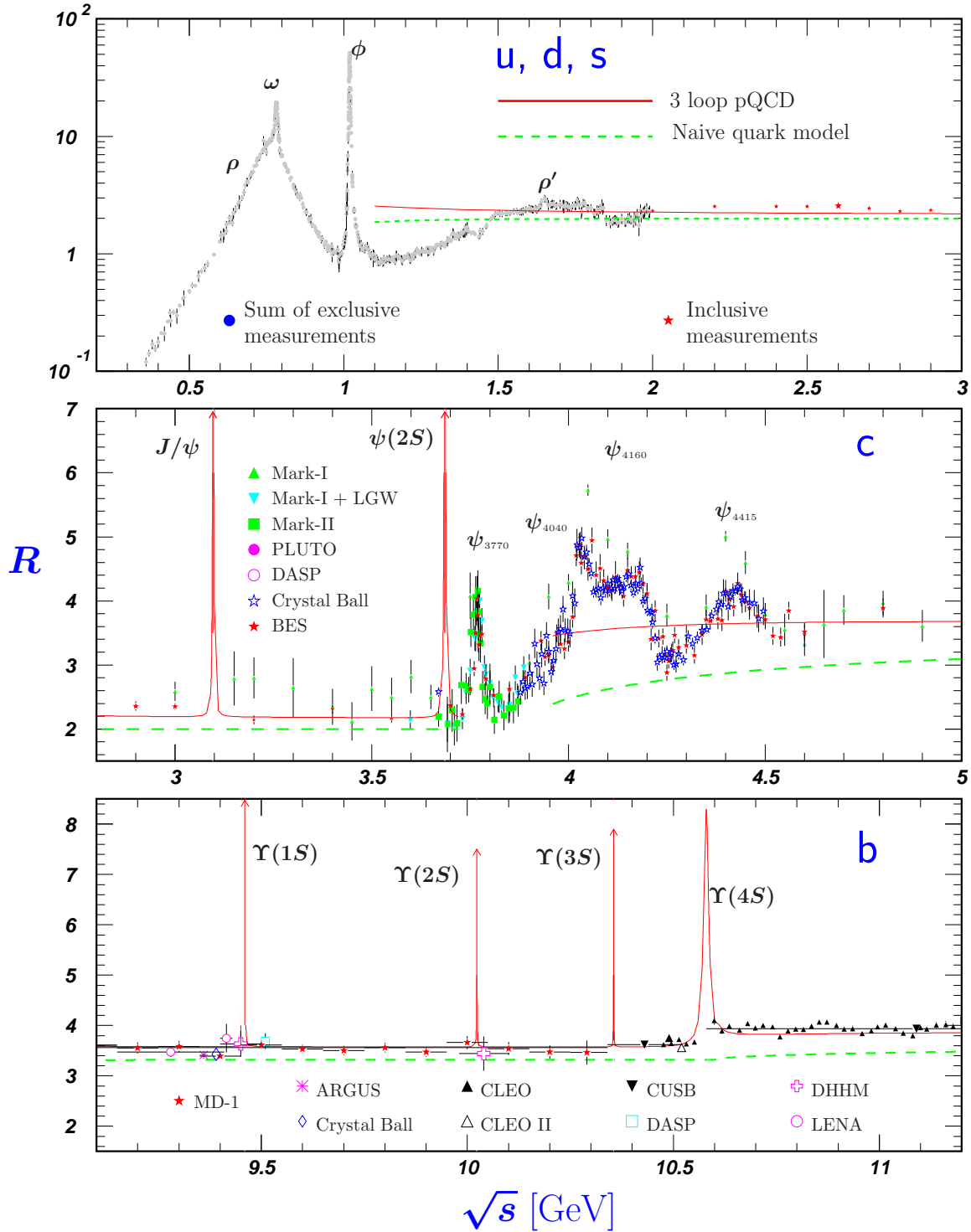
## References for Table 51.1:

- RPP 1992:** Phys. Rev. **D45** (1992); **RPP 1994:** Phys. Rev. **D50**, 1173 (1994); **RPP 1996:** Phys. Rev. **D54**, 1 (1996); **RPP 1998:** Eur. Phys. J. **C3**, 1 (1998); **RPP 2000:** Eur. Phys. J. **C15**, 1 (2000); **RPP 2002:** Phys. Rev. **D66**, 010001 (2002); **RPP 2004:** Phys. Lett. **B592**, 1 (2004); **RPP 2006:** J. Phys. **G33**, 1 (2006); **RPP 2008:** Phys. Lett. **B667**, 1 (2008); **RPP 2010:** J. Phys. **G37**, 075021 (2010); **RPP 2012:** Phys. Rev. D 86,010001(2012) and references therein.
- R.** Marshall, Rept. on Prog. in Phys. **52**, 1329 (1989). **A.** De Angelis, J. Phys. **G19**, 1233 (1993) and references therein.
- ALEPH:** D. Buskulic *et al.*: Phys. Lett. **B295**, 396 (1992); Z. Phys. **C64**, 361 (1994); **C69**, 15 (1996); **C69**, 379 (1996); **C73**, 409 (1997); and R. Barate *et al.*: Z. Phys. **C74**, 451 (1997); Phys. Reports **294**, 1 (1998); Eur. Phys. J. **C5**, 205 (1998); **C16**, 597 (2000); **C16**, 613 (2000); and A. Heister *et al.*: Phys. Lett. **B526**, 34 (2002); **B528**, 19 (2002).
- ARGUS:** H. Albrecht *et al.*: Phys. Lett. **230B**, 169 (1989); Z. Phys. **C44**, 547 (1989); **C46**, 15 (1990); **C54**, 1 (1992); **C58**, 199 (1993); **C61**, 1 (1994); Phys. Rep. **276**, 223 (1996).
- BaBar:** B. Aubert *et al.*: Phys. Rev. Lett. **87**, 162002 (2001); Phys. Rev. **D65**, 091104 (2002); J.P. Lees *et al.*: Phys. Rev. **D88**, 032011 (2013).
- Belle:** K. Abe *et al.*, Phys. Rev. Lett. **88**, 052001 (2002); and R. Seuster *et al.*, Phys. Rev. **D73**, 032002 (2006).
- CELLO:** H.J. Behrend *et al.*: Z. Phys. **C46**, 397 (1990); **C47**, 1 (1990).
- CLEO:** D. Bortoletto *et al.*, Phys. Rev. **D37**, 1719 (1988); erratum *ibid.* **D39**, 1471 (1989); and M. Artuso *et al.*, Phys. Rev. **D70**, 112001 (2004).
- Crystal Ball:** Ch. Bieler *et al.*, Z. Phys. **C49**, 225 (1991).
- DELPHI:** P. Abreu *et al.*: Z. Phys. **C57**, 181 (1993); **C59**, 533 (1993); **C61**, 407 (1994); **C65**, 587 (1995); **C67**, 543 (1995); **C68**, 353 (1995); **C73**, 61 (1996); Nucl. Phys. **B444**, 3 (1995); Phys. Lett. **B341**, 109 (1994); **B345**, 598 (1995); **B361**, 207 (1995); **B372**, 172 (1996); **B379**, 309 (1996); **B416**, 233 (1998); **B449**, 364 (1999); **B475**, 429 (2000); Eur. Phys. J. **C6**, 19 (1999); **C5**, 585 (1998); **C18**, 203 (2000); and J. Abdallah *et al.*, Phys. Lett. **B569**, 129 (2003); Phys. Lett. **B576**, 29 (2003); Eur. Phys. J. **C44**, 299 (2005); and W. Adam *et al.*: Z. Phys. **C69**, 561 (1996); **C70**, 371 (1996).
- HRS:** S. Abachi *et al.*, Phys. Rev. Lett. **57**, 1990 (1986); and M. Derrick *et al.*, Phys. Rev. **D35**, 2639 (1987).
- L3:** M. Acciarri *et al.*: Phys. Lett. **B328**, 223 (1994); **B345**, 589 (1995); **B371**, 126 (1996); **B371**, 137 (1996); **B393**, 465 (1997); **B404**, 390 (1997); **B407**, 351 (1997); **B407**, 389 (1997), erratum *ibid.* **B427**, 409 (1998); **B453**, 94 (1999); **B479**, 79 (2000).
- MARK II:** H. Schellman *et al.*, Phys. Rev. **D31**, 3013 (1985); and G. Wormser *et al.*, Phys. Rev. Lett. **61**, 1057 (1988).
- JADE:** W. Bartel *et al.*, Z. Phys. **C20**, 187 (1983); and D.D. Pietzl *et al.*, Z. Phys. **C46**, 1 (1990).
- OPAL:** R. Akers *et al.*: Z. Phys. **C63**, 181 (1994); **C66**, 555 (1995); **C67**, 389 (1995); **C68**, 1 (1995); and G. Alexander *et al.*: Phys. Lett. **B358**, 162 (1995); Z. Phys. **C70**, 197 (1996); **C72**, 1 (1996); **C72**, 191 (1996); **C73**, 569 (1997); **C73**, 587 (1997); Phys. Lett. **B370**, 185 (1996); and K. Ackerstaff *et al.*: Z. Phys. **C75**, 192 (1997); Phys. Lett. **B412**, 210 (1997); Eur. Phys. J. **C1**, 439 (1998); **C4**, 19 (1998); **C5**, 1 (1998); **C5**, 411 (1998); and G. Abbiendi *et al.*: Eur. Phys. J. **C16**, 185 (2000); **C17**, 373 (2000).
- PLUTO:** Ch. Berger *et al.*, Phys. Lett. **104B**, 79 (1981).
- SLD:** K. Abe, Phys. Rev. **D59**, 052001 (1999); Phys. Rev. **D69**, 072003 (2004).
- TASSO:** H. Aihara *et al.*, Z. Phys. **C27**, 27 (1985).
- TPC:** H. Aihara *et al.*, Phys. Rev. Lett. **53**, 2378 (1984).

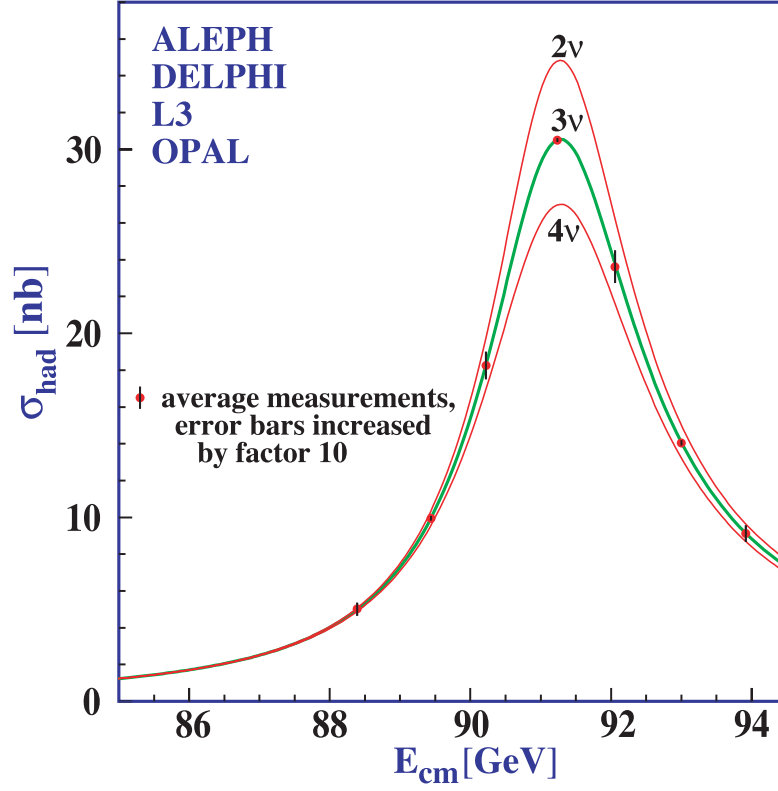
$\sigma$  and  $R$  in  $e^+e^-$  Collisions

**Figure 51.5:** World data on the total cross section of  $e^+e^- \rightarrow \text{hadrons}$  and the ratio  $R(s) = \sigma(e^+e^- \rightarrow \text{hadrons}, s) / \sigma(e^+e^- \rightarrow \mu^+\mu^-, s)$ .  $\sigma(e^+e^- \rightarrow \text{hadrons}, s)$  is the experimental cross section corrected for initial state radiation and electron-positron vertex loops,  $\sigma(e^+e^- \rightarrow \mu^+\mu^-, s) = 4\pi\alpha^2(s)/3s$ . Data errors are total below 2 GeV and statistical above 2 GeV. The curves are an educative guide: the broken one (green) is a naive quark-parton model prediction, and the solid one (red) is 3-loop pQCD prediction (see “Quantum Chromodynamics” section of this Review, Eq. (9.7) or, for more details, K. G. Chetyrkin *et al.*, Nucl. Phys. **B586**, 56 (2000) (Erratum *ibid.* **B634**, 413 (2002))). Breit-Wigner parameterizations of  $J/\psi$ ,  $\psi(2S)$ , and  $\Upsilon(nS)$ ,  $n = 1, 2, 3, 4$  are also shown. The full list of references to the original data and the details of the  $R$  ratio extraction from them can be found in [arXiv:hep-ph/0312114]. Corresponding computer-readable data files are available at <http://pdg.lbl.gov/current/xsect/>. (Courtesy of the COMPAS (Protvino) and HEPDATA (Durham) Groups, August 2015. Corrections by P. Janot (CERN) and M. Schmitt (Northwestern U.)

### $R$ in Light-Flavor, Charm, and Beauty Threshold Regions



**Figure 51.6:**  $R$  in the light-flavor, charm, and beauty threshold regions. Data errors are total below 2 GeV and statistical above 2 GeV. The curves are the same as in Fig. 51.5. **Note:** CLEO data above  $\Upsilon(4S)$  were not fully corrected for radiative effects, and we retain them on the plot only for illustrative purposes with a normalization factor of 0.8. The full list of references to the original data and the details of the  $R$  ratio extraction from them can be found in [arXiv:hep-ph/0312114]. The computer-readable data are available at <http://pdg.lbl.gov/current/xsect/>. (Courtesy of the COMPAS (Protvino) and HEPDATA (Durham) Groups, August 2015.)

Annihilation Cross Section Near  $M_Z$ 

**Figure 51.7:** Combined data from the ALEPH, DELPHI, L3, and OPAL Collaborations for the cross section in  $e^+e^-$  annihilation into hadronic final states as a function of the center-of-mass energy near the Z pole. The curves show the predictions of the Standard Model with two, three, and four species of light neutrinos. The asymmetry of the curve is produced by initial-state radiation. Note that the error bars have been increased by a factor ten for display purposes. References:

**ALEPH:** R. Barate *et al.*, Eur. Phys. J. **C14**, 1 (2000).

**DELPHI:** P. Abreu *et al.*, Eur. Phys. J. **C16**, 371 (2000).

**L3:** M. Acciarri *et al.*, Eur. Phys. J. **C16**, 1 (2000).

**OPAL:** G. Abbiendi *et al.*, Eur. Phys. J. **C19**, 587 (2001).

**Combination:** The ALEPH, DELPHI, L3, OPAL, SLD Collaborations, the LEP Electroweak Working Group, and the SLD Electroweak and Heavy Flavor Groups, Phys. Rept. **427**, 257 (2006) [[arXiv:hep-ex/0509008](#)].

(Courtesy of M. Grünewald and the LEP Electroweak Working Group, 2007)

## Total Hadronic Cross Sections

(Updated June 2016, COMPAS group, IHEP, Protvino)

This updated version of the total hadronic cross sections review is based on the first half of 2015 update of the database for total cross section and the ratio of the real-to-imaginary parts of the forward elastic scattering hadronic amplitudes. New data on total  $pp$  collisions cross sections from CERN-LHC-TOTEM [1–4] and CERN-LHC-ATLAS [5] were added.

We use a procedure for ranking models as described in [6] to identify the safest parameterizations for extrapolations. Incidentally, the models giving the best fit of accelerator data also reproduce the experimental cosmic ray nucleon–nucleon data extracted from nucleon–air data with no need of any extra phenomenological corrections to the data.

The statement in [6] that the models with universal (across of collision initial states)  $B \log^2(s/s_0)$  asymptotic term work much better than the models with  $B \log(s/s_0)$  or  $B(s/s_0)^\Delta$  terms was confirmed in [7–11] based on matching traditional asymptotic parameterizations with low energy data in different ways. However in these references the scale parameter  $s_0$  was still claimed to be dependent on the colliding particles as it should be for the asymptotic form of parameterizations constructed by the Regge-Gribov phenomenology prescriptions.

The possibility of the universal  $\log^2(s/s_0)$  rise of the hadronic total cross sections for different colliding particles was first pointed out by W. Heisenberg [12–13] and discussed many times (see for example [14] and more recent [15–17], and references therein). In [16] the universality of the asymptotic total collision cross sections has been advocated for hadron-nucleus collisions. In [17] additional indications to the universal asymptotic high-energy behavior for hadronic total collision cross sections in form  $B \log^2(s/s_0)$  were obtained from lattice QCD.

In this review we use HPR<sub>1</sub>R<sub>2</sub> model of highest COMPETE-rank modified (as in 2012 version) to save the universality of the rising part in new form that explicitly includes dependence of the  $s_0$  and  $B$  on the initial state mass parameters and the new scale parameter  $M$ .

$$\sigma^{a\bar{b}} = H \log^2 \left( \frac{s}{s_M^{ab}} \right) + P^{ab} + R_1^{ab} \left( \frac{s}{s_M^{ab}} \right)^{-\eta_1} \pm R_2^{ab} \left( \frac{s}{s_M^{ab}} \right)^{-\eta_2};$$

$$\rho^{a\bar{b}} = \frac{1}{\sigma^{a\bar{b}}} \left[ \pi H \log \left( \frac{s}{s_M^{ab}} \right) - R_1^{ab} \left( \frac{s}{s_M^{ab}} \right)^{-\eta_1} \tan \left( \frac{\eta_1 \pi}{2} \right) \pm R_2^{ab} \left( \frac{s}{s_M^{ab}} \right)^{-\eta_2} \cot \left( \frac{\eta_2 \pi}{2} \right) \right],$$

where upper signs in formulas are for particles and lower signs for antiparticles. The adjustable parameters are as follows:

$H = \pi \frac{(hc)^2}{M^2}$  in mb, where notation  $H^\S$  is after Heisenberg(1952,1975);

$P^{ab}$  in mb, are Pomeranchuk's(1958) constant terms;

$R_i^{ab}$  in mb are the intensities of the effective secondary Regge pole contributions named after Regge-Gribov(1961);

$s, s_M^{ab} = (m_a + m_b + M)^2$  are in  $\text{GeV}^2$ ;

$m_a, m_b, (m_{\gamma^*} = m_{\rho(770)})$  are the masses of initial state particles, and  $M$  – the mass parameter defining the rate of universal rise of the cross sections are all in  $\text{GeV}$ . Parameters  $M, \eta_1$  and  $\eta_2$  are universal for all collisions considered.

Exact factorization hypothesis was used for both  $H \log^2(\frac{s}{s_M^{ab}})$  and  $P^{ab}$  to extend the universal rise of the total hadronic cross sections to the  $\gamma(p, d) \rightarrow \text{hadrons}$  and  $\gamma\gamma \rightarrow \text{hadrons}$  collisions. This results in one additional adjustable parameter  $\delta$  with substitutions:

$$\pi H \log^2 \left( \frac{s}{s_M^{\gamma(p,d)}} \right) + P^{\gamma(p,d)} \Rightarrow \delta \left[ \pi(1, \lambda) H \log^2 \left( \frac{s}{s_M^{\gamma(p,d)}} \right) + P^{p(p,d)} \right];$$

$$\pi H \log^2 \left( \frac{s}{s_M^{\gamma\gamma}} \right) + P^{\gamma\gamma} \Rightarrow \delta^2 \left[ \pi H \log^2 \left( \frac{s}{s_M^{\gamma\gamma}} \right) + P^{pp} \right].$$

These parameterizations were used for simultaneous fit with **35** adjustable parameters to the data on collisions:

$$(\bar{p}, p) (p, n, d); \quad \Sigma^- p; \quad \pi^\mp (p, n, d); \quad K^\mp (p, n, d); \quad \gamma p; \quad \gamma \gamma; \quad \gamma d.$$

The results of the fits are presented in the following table and figures. In the table, two values of the fit quality indicator  $\text{FQ} = \chi^2/(\text{Npt} - 35)$  are reported in the last element of the first row, where **Npt** is the number of data points in corresponding sample.  $\text{FQ}_{\text{INT}}$  calculated with “internal” parameter values of machine precision (16 digits) and  $\text{FQ}_{\text{EXT}}$  calculated with rounded parameter values as displayed in the table in accordance with PDG rules (Section 5.3 of J. Beringer *et al.*, (Particle Data Group), Phys. Rev. **D86**, 010001 (2012)), recent metrology recommendations [18] and rules for safe uniform rounding of correlated data [19]. The uniformity of the quality of data description across different collisions is shown in the last two columns of the table; **npt** is the number of data points in a subsample and  $\chi^2/\text{npt}$  is the contribution of the subsample to the global  $\chi^2$  reduced to **npt**.

<sup>\S</sup>For collisions with deuteron target  $H_d = \lambda H$  where dimensionless parameter  $\lambda$  is introduced to test the universality of the Heisenberg rise for particle–nuclear and nuclear–nuclear collisions.

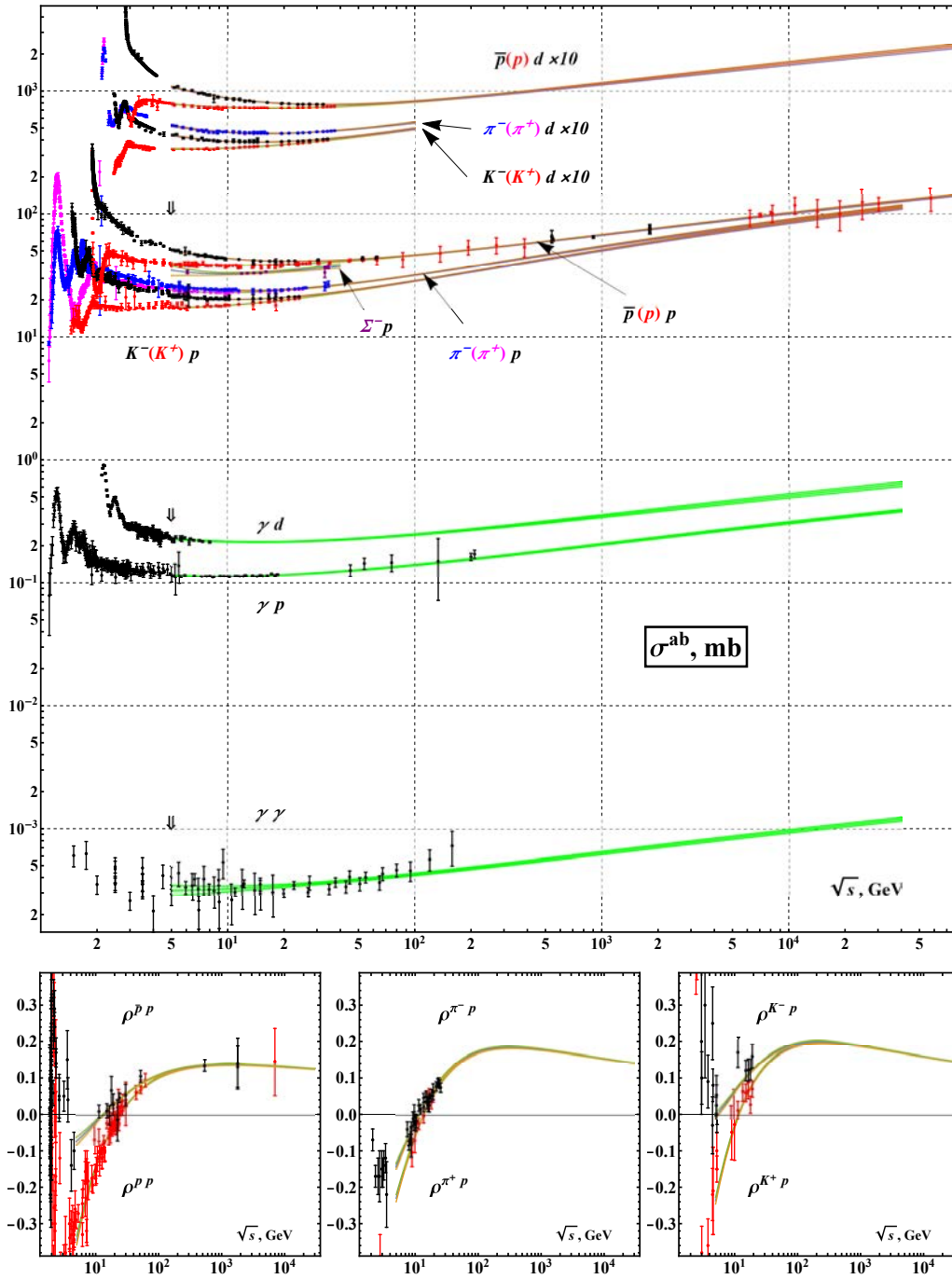
HPR <sub>1</sub> R <sub>2</sub> at $\sqrt{s} \geq 5\text{GeV}$	M=2.1206 $\pm$ 0.0094 [GeV]		H=0.2720 $\pm$ 0.0024 [mb]		FQ <sub>INT</sub> = 0.96
	$\eta_1 = 0.4473 \pm 0.0077$		$\eta_2 = 0.5486 \pm 0.0049$		FQ <sub>EXT</sub> = 0.96
	$\delta = (3.063 \pm 0.014) \times 10^{-3}$		$\lambda = 1.624 \pm 0.033$		
P[mb]	R <sub>1</sub> [mb]	R <sub>2</sub> [mb]	Beam/Target	Npt=1048	$\chi^2/npt$ by Groups
34.41 $\pm$ 0.13	13.07 $\pm$ 0.17	7.394 $\pm$ 0.081	$\bar{p}(p)/p$	258	1.14
34.71 $\pm$ 0.17	12.52 $\pm$ 0.34	6.66 $\pm$ 0.15	$\bar{p}(p)/n$	67	0.48
34.7 $\pm$ 1.3	-46. $\pm$ 18.	-48. $\pm$ 18.	$\Sigma^-/p$	9	0.37
18.75 $\pm$ 0.11	9.56 $\pm$ 0.15	1.767 $\pm$ 0.030	$\pi^\mp/p$	183	1.02
16.36 $\pm$ 0.09	4.29 $\pm$ 0.13	3.408 $\pm$ 0.044	$K^\mp/p$	121	0.82
16.31 $\pm$ 0.10	3.70 $\pm$ 0.19	1.826 $\pm$ 0.068	$K^\mp/n$	64	0.58
	0.0139 $\pm$ 0.0011		$\gamma/p$	41	0.62
	(-4. $\pm$ 17.) $\times 10^{-6}$		$\gamma/\gamma$	37	0.75
	0.0370 $\pm$ 0.0019		$\gamma/d$	13	0.9
64.45 $\pm$ 0.32	29.66 $\pm$ 0.39	14.94 $\pm$ 0.18	$\bar{p}(p)/d$	85	1.52
36.65 $\pm$ 0.26	18.75 $\pm$ 0.36	0.341 $\pm$ 0.091	$\pi^\mp/d$	92	0.72
32.06 $\pm$ 0.19	7.70 $\pm$ 0.31	5.616 $\pm$ 0.082	$K^\mp/d$	78	0.79

To construct the parameter scatter region we follow Section 39.4.2.2 of J. Beringer *et al.* (Particle Data Group), Phys. Rev. **D86**, 010001 (2012) and recent metrology JCGM 101:2008 recommendations and produce the direct Monte Carlo propagation of uncertainties from experimental data to the uncertainties of the best fit parameters. To do this we interpret the whole input data sample as statistically independent sample with total experimental uncertainty at each experimental data point being a Gaussian standard deviation. This technical assumption allows us to generate MC sampling of experimental data and to obtain at each MC trial new “biased” best fit parameters belonging to scatter region of the initial best fit parameters values. These biased best fit parameters constitute the MC-samples of cardinalities  $|MC_{cut}|$  at each  $\sqrt{s}$  cutoff and are the basis for construction of three 35-dimensional empirical parameter distributions.

In paper [19] the asymptotic bounds (Froissart, Martin) on the possible rise of the total collision cross sections in the form  $\log^2(s/s_0)$  was questioned in favour of possible faster rising forms. It was supported by the fits presented in [18] where the form  $\log^c(s/s_0)$  with adjustable  $c$  was tested on  $(\bar{p})pp$  data only and it was claimed that values of  $c$  obtained in number of different fits are statistically compatible with  $c \in [2.2, 2.4]$ .

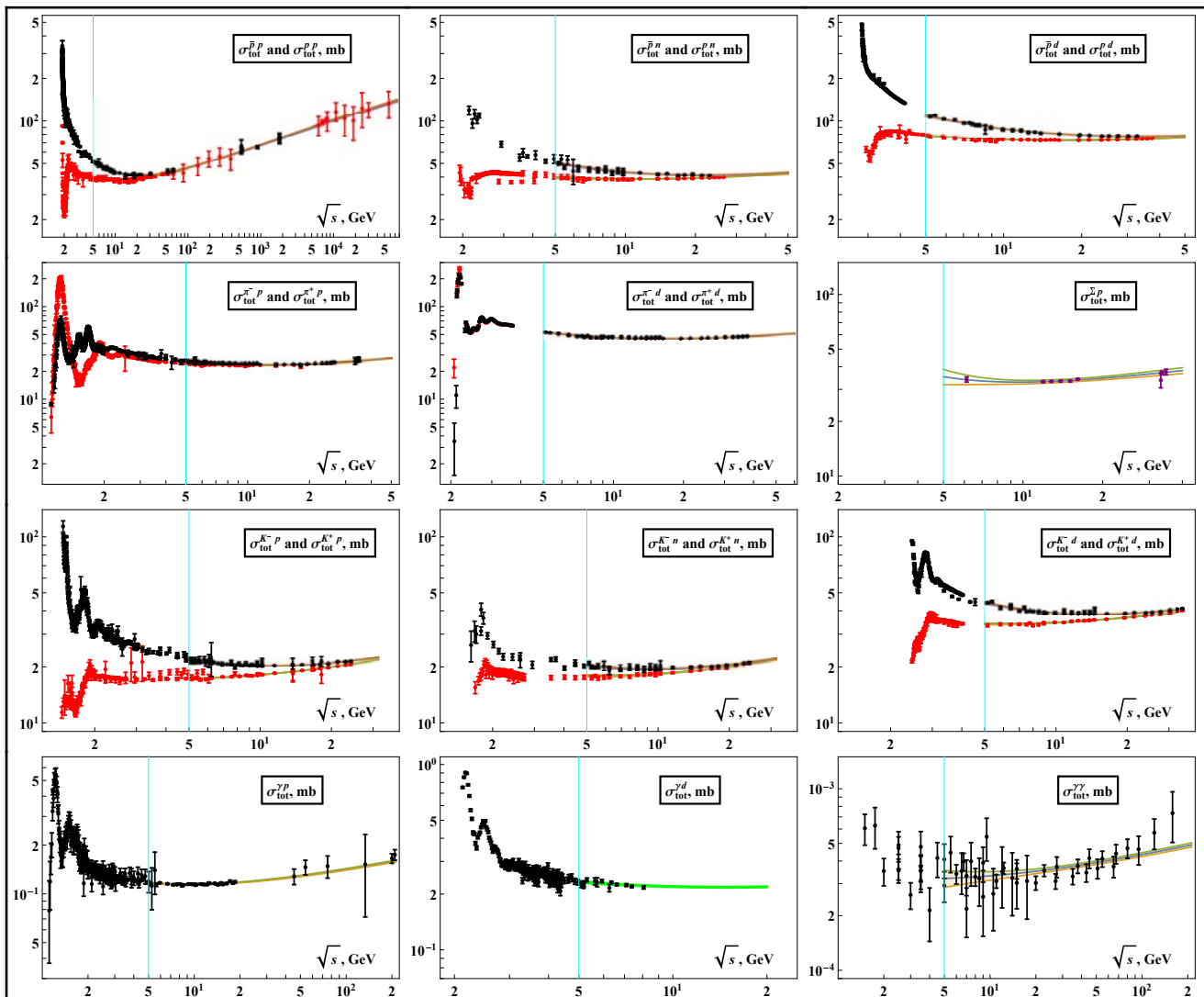
We have performed our global fit with adjustable  $c$  to the total cross sections and available  $p$ -parameters (as of August 2015) including TOTEM data point at 8 TeV [20]. For this fit we have 36 adjustable parameters. Fit was done with all data at  $\sqrt{s} \geq 5$  GeV with  $FQ = 0.87$ . We have obtained value  $c = 1.98 \pm 0.01$  (Hessian error) which is in two standard deviation lower than  $c = 2$  (exact) and possibly could be tentatively interpreted as an indication to the slower universal rising total cross sections as it was proposed 45 years ago by Cheng and Wu in the form  $\log^2\left(\frac{(s/s_0)^a}{\log^2(s/s_0)}\right)$  in their seminal paper [21]. However, to notice this difference much experimental, theoretical, and modelling work has to be done.

In conclusion, the Heisenberg prediction of the universal  $\log^2(s/s_M)$  form of asymptotic rise of the hadronic collision total cross sections is still actual and should be tested in all aspects at available colliders operating with  $(\bar{p}, p, nuclei)$  beams and in experiments with cosmic rays.



**Figure 51.8:** Summary of  $h^\mp p \rightarrow \text{anything}$ ,  $\gamma p \rightarrow \text{hadrons}$ ,  $\gamma \gamma \rightarrow \text{hadrons}$  total cross sections  $\sigma^{ab}$  in mb and  $\rho^{h^\mp p}$  the ratio of real to imaginary parts of the forward hadronic amplitudes. Also for qualitative comparison of the uniformity of data description by HPR<sub>1</sub>R<sub>2</sub>-model across the different collisions and observables. The uncertainties for the experimental data points include both the statistical and systematic errors. Curves, corresponding to fit above 5 GeV cut, are plotted with error bands calculated with parameter covariance matrix constructed on MC-propagated vectors from 95% quantile of the empirical distribution.





**Figure 51.9:** Summary of all total collision cross sections jointly fitted with available hadronic  $\rho$  parameter data. Corresponding computer-readable data files may be found at <http://pdg.lbl.gov/xsect/contents.html>

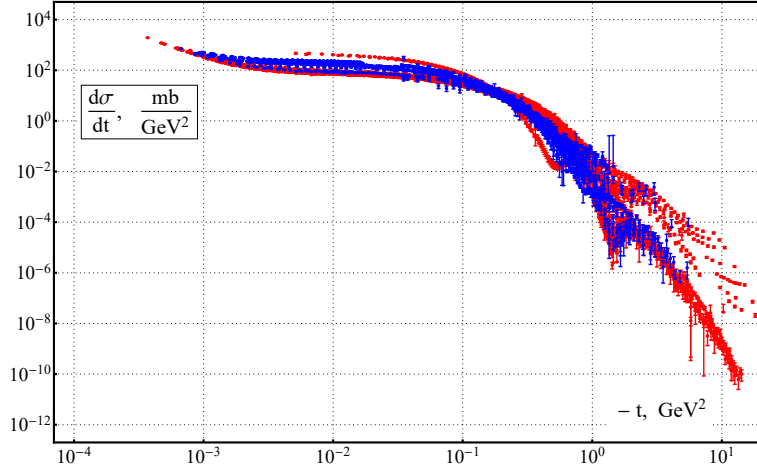
## References

- [1] G. Antchev *et al.*, Nucl. Phys. **B899**, 527 (2013).
- [2] G. Antchev *et al.*, Phys. Rev. Lett. **111**, 012001 (2013).
- [3] G. Antchev *et al.*, Europhys. Lett. **101**, 21004 (2013).
- [4] G. Antchev *et al.*, Europhys. Lett. **101**, 21002 (2013).
- [5] G. Aad *et al.*, Nucl. Phys. **B889**, 486 (2014).
- [6] J.R. Cudell *et al.*: Phys. Rev. **D65**, 074024 (2002); Phys. Rev. Lett. **89**, 201801 (2002).
- [7] K. Igi and M. Ishida: Phys. Rev. **D66**, 034023 (2002); Phys. Lett. **B622**, 286 (2005).
- [8] M.M. Block and F. Halzen: Phys. Rev. **D72**, 036006 (2005); Phys. Rev. **D72**, 039902 (2005).
- [9] M. Ishida and K. Igi, Prog. Theor. Phys. Suppl. **187**, 297 (2011).
- [10] M. Ishida and V. Barger, Phys. Rev. **D84**, 014027 (2011).
- [11] F. Halzen *et al.*, Phys. Rev. **D85**, 074020 (2012).
- [12] W. Heisenberg, Z. Phys. **133**, 65 (1952).
- [13] W. Heisenberg, *Fourteenth Int. Cosm. Ray Conference*, Vol.11, München 1975, 3461-3474;
- [14] S.S. Gershtein and A.A. Logunov, Sov. J. Nucl. Phys. **39**, 960 (1984).
- [15] E. Iancu and R. Venugopalan, R.C. Hwa (ed.) *et al.*, [hep-ph/0303204].
- [16] L. Frankfurt, M. Strikman, and M. Zhalov, Phys. Lett. **B616**, 59 (2005).
- [17] M. Giordano, E. Meggiolaro, and N. Moretti, JHEP **1209**, 031 (2012).
- [18] JCGM 100:2008, JCGM 101:2008, JCGM 104:2009, JCGM 102:2011 via [www.bipm.org/en/publications/guides/gum.html](http://www.bipm.org/en/publications/guides/gum.html)
- [19] V.V. Ezhela, Data Science Journal **6**, PS676 (2007).
- [20] Y.I. Azimov, Phys. Rev. **D84**, 056012 (2011).
- [21] D.A. Fagundes, M.J. Menon, and P.V.R.G. Silva, J. Phys. **G40**, 065005 (2013).
- [22] H. Cheng and T.T. Wu, Phys. Rev. Lett. **24**, 1456 (1970).

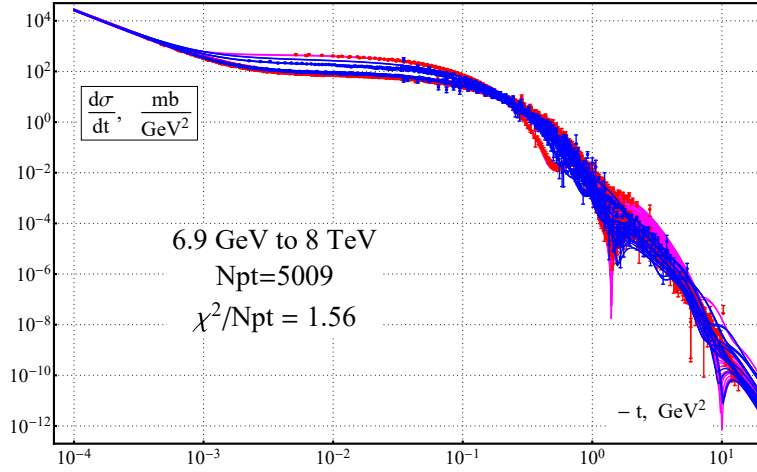
## High Energy Elastic $\bar{p}p$ and $pp$ Differential Cross Sections

(Updated June 2016, COMPAS group, IHEP, Protvino)

Using new results from FNAL-COLLIDER-D0 experiment in  $\bar{p}p$  elastic collisions at  $\sqrt{s} = 1.96$  TeV [1], CERN-LHC-TOTEM, CERN-LHC-ATLAS experiments in  $pp$  elastic collisions at  $\sqrt{s} = 7, 8$  TeV [2–3] and PAO experiment in proton-air collisions at 57 TeV [4] the amplitudes of the elastic  $\bar{p}p$  and  $pp$  collisions are investigated in a most broad region in  $\sqrt{s}$  and  $t$  via three observables  $d\sigma/dt(s, t)$ ,  $\sigma^{tot}(s)$ , and  $\rho(s)$ . The summary of the database for  $d\sigma/dt(s, t)$  is presented in Figure 51.10, where projection of the  $d\sigma/dt(\sqrt{s}, t)$  to the  $(d\sigma/dt, -t)$  plane orthogonal to the  $\sqrt{s}$  axis is displayed.



**Figure 51.10:** Cumulative plots of data on  $d\sigma/dt$  for  $\bar{p}p$  (blue) and  $pp$  (red) elastic collisions at  $\sqrt{s} \geq 2.99$  GeV. Number of data points  $N_{pt} = 6629$

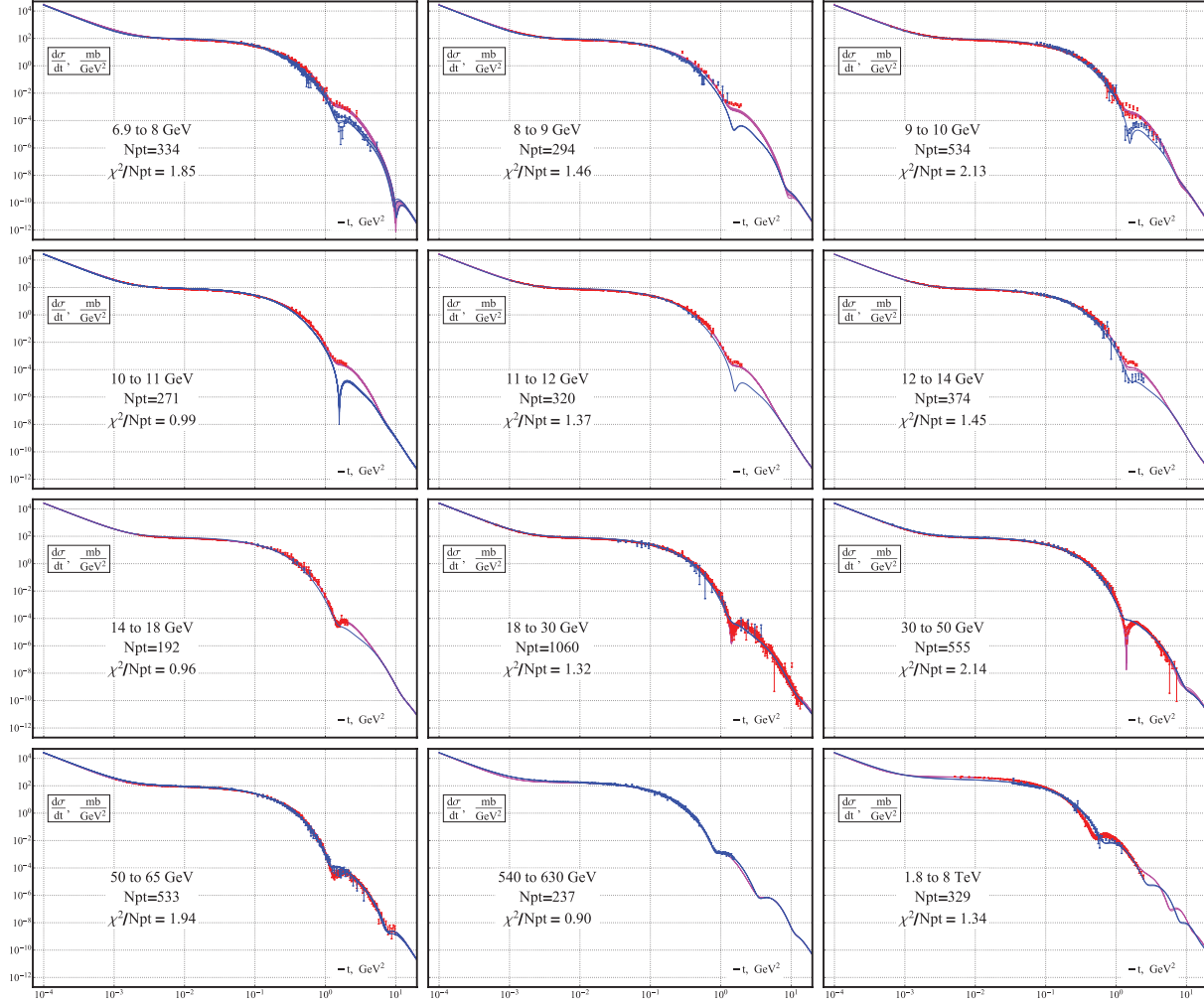


**Figure 51.11:** Cumulative plots of data on  $d\sigma/dt$  and model description for  $\bar{p}p$  (blue,blue) and  $pp$  (red,magenta) elastic collisions at  $\sqrt{s} \geq 6.9$  GeV.

All characteristic features of the  $d\sigma/dt(\sqrt{s}, t)$  behavior in  $-t$  and  $\sqrt{s}$  are clearly seen:

- The energy-dependent Coulomb-Nuclear Interference (CNI) effects at small  $-t$ ;
- Diffractive peaks with crossover effect at  $-t \approx 0.16$  GeV<sup>2</sup> for particle-antiparticle data at same energies;
- The first dip/shoulder moving to the left with growing  $\sqrt{s}$ . New data on  $d\sigma/dt$  in  $\bar{p}p$  and  $pp$  elastic collisions at highest accelerator energies have challenged all previous model predictions that gave “not so bad qualitative agreement” with previously available data on  $d\sigma/dt$ . There is a need to reveal a quantitative and statistically complete picture of the data description by at least one model with most ambitious claim on the “best known description”. There are several conceptually related papers with such a claims [5–7] but with different areas of applicability and without treatment of the CNI region. Description of  $d\sigma/dt$  by our model (a variation of AGNM [7] parameterization) at  $\sqrt{s} \geq 6.9$  GeV is displayed on Figure 51.11.

Historically the most complete compilations on  $d\sigma/dt$  data expressed in Mandelstam variables  $\sqrt{s}$  and  $t$  were published in Landolt-Börnstein volumes (now available in digital form) up to 1981 [8]. Updated (in high energy part) analogous CLM-compilation [9] (available in computer readable form) was compiled with help of HEPDATA and COMPAS databases and released in 2006. In our fits we use the CLM-compilation with minor corrections, filled detected gaps, and updated with new data published up to August 2015. We performed simultaneous fits to the sample of data on  $d\sigma/dt(s, t)$ ,  $\sigma^{tot}(s)$ , and  $\rho(s)$  in  $\bar{p}p$  and  $pp$  collisions at  $6.9 \text{ GeV} \leq \sqrt{s} \leq 8 \text{ TeV}$  and all available  $t$ . Overall fit quality  $FQ = \chi^2(N_{pt})/(5266 - 37) = 1.60$ , which is unreliable for our number of degrees of freedom. Removing contributions to  $\chi^2(d\sigma/dt)$  from  $N_{out} = 277$  points with  $\chi^2(point) > (2.4)^2$  (of  $2.4 \times \text{standard deviation (std)}$ ) – randomly scattered outliers) we have  $\chi^2(N_{rest})/(N_{rest}) = 1.02$ , where  $N_{rest} = N_{pt}(d\sigma/dt) - N_{out}$ . The uniformity level of the fit quality in different intervals of  $\sqrt{s}$  is shown on Figure 51.12.



**Figure 51.12:** All 12 energy intervals are non-overlapping and cover all data points. All panels have axis labelled as in Figure 51.11. All data sets corresponding to the same energy have model curve drawn. Some panels have no red or blue data points. In such a cases we add the curve as prediction of the model.

To reveal a more complete picture of phenomenological description of the elastic scattering data we originally selected the most flexible model (43 parameters) from [7] with most broad claimed area of applicability. It turns out that without partial removal of some data (made in [7]) the claimed solution cannot be used even as starting point for adjustments, most probably, because of over-rounded published parameter values, or misprints in parameter tables, or strong parameter correlations. Moreover, numerous fits with different starting points failed to find any locally stable solution with physically reasonable adjustable parameter values. We obtain a stable solution only with addition of data in CNI region and with slightly modified parameterization to reduce the number of adjustable parameters from 43 to 37. Our expressions for observables  $\sigma_{\pm}^{tot}(s)$ ,  $\rho_{\pm}(s)$ , and  $d\sigma_{\pm}/dt(s, t)$  with sign “+” for  $pp$  and sign “-” for  $\bar{p}p$  collisions are constructed (in notations of [7]) using corresponding scattering amplitudes: nuclear  $T_{\pm}(s, t)$  and Coulomb  $T_{\pm}^c(s, t)$  both in  $\text{mb} \cdot \text{GeV}^2$  as follows:

$$\sigma_{\pm}^{tot}(s) = \frac{\text{Im } T_{\pm}(s, 0)}{\sqrt{s(s - 4m_p^2)}}, \quad \rho_{\pm}(s) = \frac{\text{Re } T_{\pm}(s, 0)}{\text{Im } T_{\pm}(s, 0)}, \quad \frac{d\sigma_{\pm}}{dt}(s, t) = \frac{|T_{\pm}(s, t) + T_{\pm}^c(s, t)|^2}{16\pi(\hbar c)^2 s(s - 4m_p^2)},$$

where constants:  $m_p$  stands for proton mass, and  $(\hbar c)^2$  for  $\text{mb-to-GeV}^2$  conversion factor. Nuclear amplitudes  $T_{\pm}(s, t)$  are linearly combined crossing even  $F_+(\hat{s}, t)$  and crossing odd  $F_-(\hat{s}, t)$  functions.

$$T_{\pm}(s, t) = [F_+(\hat{s}, t) \pm F_-(\hat{s}, t)],$$

$$F_+(\hat{s}, t) = F_+^H(\hat{s}, t) + F_+^P(\hat{s}, t) + F_+^{PP}(\hat{s}, t) + F_+^R(\hat{s}, t) + F_+^{RP}(\hat{s}, t) + N_+(s, t),$$

$$F_-(\hat{s}, t) = F_-^{MO}(\hat{s}, t) + F_-^O(\hat{s}, t) + F_-^{OP}(\hat{s}, t) + F_-^R(\hat{s}, t) + F_-^{RP}(\hat{s}, t) + N_-(s, t),$$

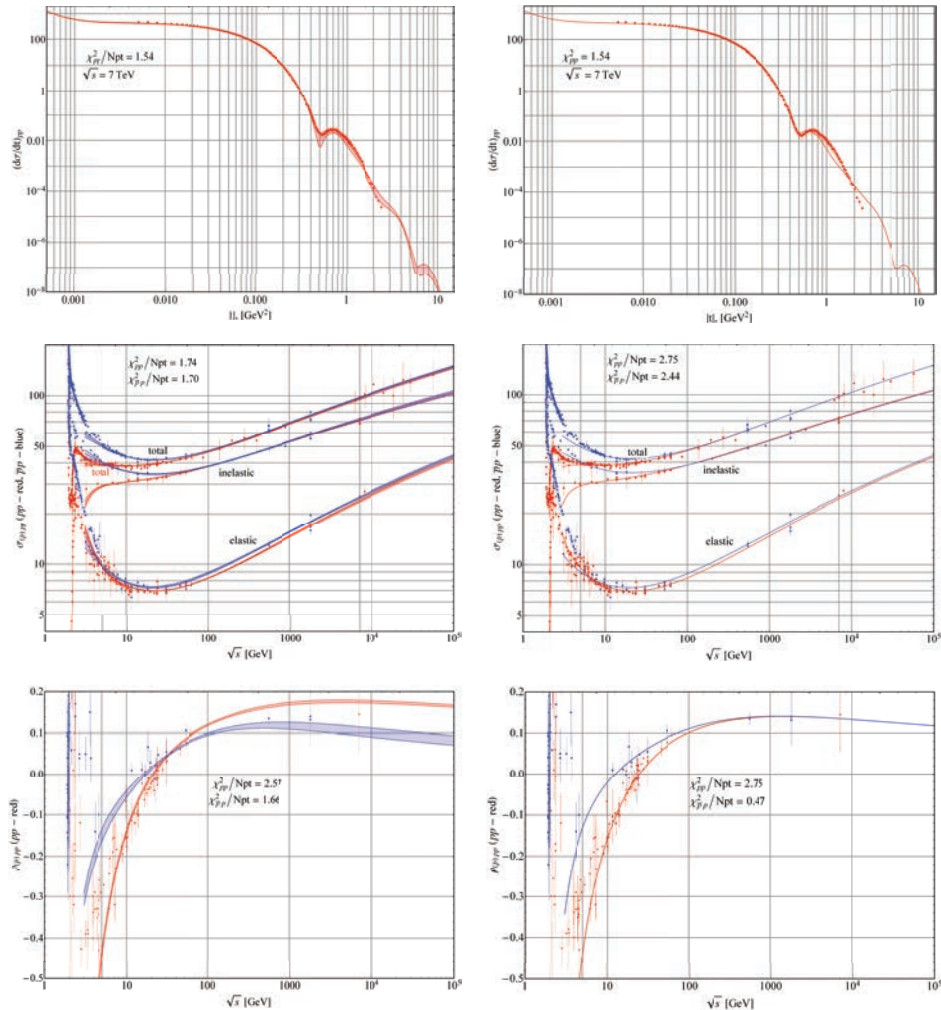
$$F_+^H(\hat{s}, t) = i\hat{s} \left\{ \begin{array}{l} H_1 \frac{2J_1(K+\tilde{\tau})}{K+\tilde{\tau}} \cdot e^{b+1t} \ln^2 \tilde{s} + \\ H_2 J_0(K+\tilde{\tau}) \cdot e^{b+2t} \ln \tilde{s} + \\ H_3 [J_0(K+\tilde{\tau}) - K+\tilde{\tau} J_1(K+\tilde{\tau})] \cdot e^{b+3t} \end{array} \right\}, \quad F_-^{MO}(\hat{s}, t) = \hat{s} \left\{ \begin{array}{l} O_1 \frac{\sin(K-\tilde{\tau})}{K-\tilde{\tau}} e^{b-1t} \cdot \ln^2 \tilde{s} + \\ O_2 \cos(K-\tilde{\tau}) e^{b-2t} \cdot \ln \tilde{s} + \\ O_3 e^{b-3t} \end{array} \right\},$$

$$\begin{aligned}
F_+^P(\hat{s}, t) &= -C_P e^{b_P t} e^{-i\frac{\pi}{2}\alpha_P(t)} (\hat{s})^{\alpha_P(t)}, & F_-^O(\hat{s}, t) &= iC_O e^{b_O t} e^{-i\frac{\pi}{2}\alpha_O(t)} (\hat{s})^{\alpha_O(t)} (1 + A_O t), \\
F_+^{PP}(\hat{s}, t) &= -\frac{C_{PP}}{\ln \hat{s}} e^{b_{PP} t} e^{-i\frac{\pi}{2}\alpha_{PP}(t)} (\hat{s})^{\alpha_{PP}(t)}, & F_-^{OP}(\hat{s}, t) &= i\frac{C_{OP}}{\ln \hat{s}} e^{b_{OP} t} e^{-i\frac{\pi}{2}\alpha_{OP}(t)} (\hat{s})^{\alpha_{OP}(t)}, \\
F_{\pm}^{RP}(\hat{s}, t) &= \frac{tC_{RP}^{\pm}}{\ln \hat{s}} e^{b_{RP}^{\pm} t} e^{-i\frac{\pi}{2}\alpha_{RP}^{\pm}(t)} (\hat{s})^{\alpha_{RP}^{\pm}(t)}, & F_{\pm}^R(\hat{s}, t) &= \mp C_R^{\pm} e^{b_R^{\pm} t} e^{-i\frac{\pi}{2}\alpha_R^{\pm}(t)} (\hat{s})^{\alpha_R^{\pm}(t)}, \\
N_{\pm}(s, t) &= -i^{\frac{1\pm 1}{2}} \cdot \hat{s} \cdot N_{\pm} \cdot (\ln \hat{s}) \frac{t}{t_0} \cdot (1 - t/t_{\pm})^{-5}, \\
\alpha_P(t) &= 1 + \alpha'_P \cdot t; & \alpha_R^{\pm}(t) &= \alpha_R^{\pm}(0) + \alpha_R^{\pm'} \cdot t; & \alpha_O(t) &= 1 + \alpha'_O \cdot t, \\
\alpha_{OP}(t) &= 1 + \frac{\alpha'_P \alpha'_O}{\alpha'_P + \alpha'_O} \cdot t; & \alpha_{PP}(t) &= 1 + \frac{\alpha'_P}{2} \cdot t; & \alpha_{RP}^{\pm}(t) &= \alpha_R^{\pm}(0) + \frac{\alpha_P' \alpha_R^{\pm'}}{\alpha_P' + \alpha_R^{\pm'}} \cdot t, \\
\hat{s}(s, t) &\equiv \hat{s} = (-t + 2s - 4m_p^2)/(2s_0), & s_0 &= 1 \text{ GeV}^2; & \ln(\hat{s}) &= \ln(\hat{s}) - i\frac{\pi}{2}; & \tilde{\tau} &= \sqrt{-t/t_0} \ln \hat{s}, & t_0 &= 1 \text{ GeV}^2.
\end{aligned}$$

Coulomb amplitudes are taken with dipole electric nucleon form factor

$$T_{\pm}^C(s, t) = \mp e^{[\pm i\alpha\Phi_{\pm}^{NC}(s, t)]} \cdot 8\pi(\hbar c)^2 \alpha \cdot \frac{s}{t} \cdot \left(1 - \frac{t}{\Lambda^2}\right)^{-4},$$

where:  $\Phi_{\pm}^{CN}(s, t) = \ln \left[ -\frac{t}{2} \left( B_{\pm}(s) + \frac{8}{\Lambda^2} \right) \right] + \gamma - \frac{4t}{\Lambda^2} \ln \left( -\frac{4t}{\Lambda^2} \right) - \frac{2t}{\Lambda^2}$  is the CNI phase in the R. Cahn form [10];  $\Lambda = \sqrt{0.71} \text{ GeV}$ ;  $\alpha$  – fine structure constant;  $\gamma$  – Euler constant. Instead of the traditional definition of the  $d\sigma_{\pm}/dt(s, t)$  slope function  $B_{\pm}(s) = \left[ \frac{d}{dt} \ln \left( \frac{d\sigma_{\pm}}{dt}(s, t) \right) \right]_{t=0}$ , we set  $B_{\pm}(s) = \frac{\sigma_{\pm}(s)}{4\pi(\hbar c)^2}$  to simplify calculations and to get faster minimization procedures. The odderon contribution at  $t = 0$  can be switched off by replacing all exponents  $e^{b_i t}$  by  $e^{b_i t}(1 - e^{-t/t_{min}})$  in the above expressions for all odderon related terms, where  $t_{min}$  is the minimal  $|t|$ -value for which the data exist in the  $d\sigma/dt$  database. The resulting non-intersecting  $\sigma_{\mp}(\sqrt{s})$  and  $\rho_{mp}(\sqrt{s})$  curves can be obtained without significant degradation to the fit quality [11]. The effect of switching odderon contribution on and off at  $t = 0$  is shown in Figure 51.13.



**Figure 51.13:** Comparisons of  $d\sigma/dt$  at  $\sqrt{s} = 7 \text{ TeV}$ ,  $d\sigma^{tot}$ , and  $\rho(\sqrt{s})$  data with our two variants of AGNM [7] models. Right-hand panels have odderons switched off at  $t = 0$ . Curves for integrated elastic cross sections and inelastic cross sections were obtained from integration of the elastic differential cross sections. Data from experiments are not used in the fits.

The best fit parameter values and their standard deviations for both models are listed in the tables below. The table on the left (right) is with (without) oderon contribution. Estimates of the std were obtained by the MC-propagation of the assumed Gaussian distribution for each individual data point. Despite of poor MC statistics, the obtained “propagated” covariance matrix is in good conditions and gives reasonable std estimates. The quality of the fit to the  $\sigma_{\mp}^{tot}(s)$  and  $\rho_{\mp}(s)$  data is presented in Figure 51.13. Error bands were calculated by propagation of the parameter scatter region to the scatter region of these observables.

Name	Unit	Value	$\pm Vstd$	Name	Unit	Value	$\pm Vstd$	Name	Unit	Value	$\pm Vstd$	Name	Unit	Value	$\pm Vstd$
$H_1$	mb GeV <sup>2</sup>	0.2488	0.0008	$O_1$	mb GeV <sup>2</sup>	<b>0.</b>	(fix)	$H_1$	mb GeV <sup>2</sup>	0.2525	0.0005	$O_1$	mb GeV <sup>2</sup>	<b>0.</b>	(fix)
$H_2$	mb GeV <sup>2</sup>	0.00691	0.00041	$O_2$	mb GeV <sup>2</sup>	0.576	0.026	$H_2$	mb GeV <sup>2</sup>	0.002545	0.000095	$O_2$	mb GeV <sup>2</sup>	0.576	0.004
$H_3$	mb GeV <sup>2</sup>	10.42	0.17	$O_3$	mb GeV <sup>2</sup>	-3.256	0.160	$H_3$	mb GeV <sup>2</sup>	9.51	0.05	$O_3$	mb GeV <sup>2</sup>	-3.256	0.003
$K_+$		0.3092	0.0010	$K_-$		0.1000	0.0012	$K_+$		0.3164	0.0035	$K_-$		0.1000	0.0008
$C_P$	mb GeV <sup>2</sup>	-0.0946	0.0058	$C_O$	mb GeV <sup>2</sup>	-6.91	0.27	$C_P$	mb GeV <sup>2</sup>	-0.2661	0.0040	$C_O$	mb GeV <sup>2</sup>	-6.910	0.032
$C_{PP}$	mb GeV <sup>2</sup>	159.80	1.44	$C_{OP}$	mb GeV <sup>2</sup>	53.83	1.12	$C_{PP}$	mb GeV <sup>2</sup>	171.80	0.75	$C_{OP}$	mb GeV <sup>2</sup>	53.83	0.06
$C_R^+$	mb GeV <sup>2</sup>	-30.20	1.04	$C_R^-$	mb GeV <sup>2</sup>	85.10	1.79	$C_R^+$	mb GeV <sup>2</sup>	-41.84	0.62	$C_R^-$	mb GeV <sup>2</sup>	85.10	0.06
$C_{RP}^+$	mb GeV <sup>2</sup>	-1.897	0.124	$C_{RP}^-$	mb GeV <sup>2</sup>	-48.77	2.77	$C_{RP}^+$	mb GeV <sup>2</sup>	-0.511	0.087	$C_{RP}^-$	mb GeV <sup>2</sup>	-48.7700	0.0021
$\alpha_R^+(0)$		0.651	0.009	$\alpha_R^-(0)$		0.4558	0.0060	$\alpha_R^+(0)$		0.628	0.006	$\alpha_R^-(0)$		0.4558	0.0060
$\alpha_R'^+$	GeV <sup>-2</sup>	<b>0.8</b>	(fix)	$\alpha_R'^-$	GeV <sup>-2</sup>	<b>0.8</b>	(fix)	$\alpha_R'^+$	GeV <sup>-2</sup>	<b>0.8</b>	(fix)	$\alpha_R'^-$	GeV <sup>-2</sup>	<b>0.8</b>	(fix)
$\alpha_P'$	GeV <sup>-2</sup>	0.1603	0.0052	$\alpha_O'$	GeV <sup>-2</sup>	0.680	0.033	$\alpha_P'$	GeV <sup>-2</sup>	0.2461	0.0013	$\alpha_O'$	GeV <sup>-2</sup>	0.680	0.026
$b_{+1}$	GeV <sup>-2</sup>	3.895	0.029	$b_{-1}$	GeV <sup>-2</sup>	<b>0.</b>	(fix)	$b_{+1}$	GeV <sup>-2</sup>	4.017	0.008	$b_{-1}$	GeV <sup>-2</sup>	<b>0.</b>	(fix)
$b_{+2}$	GeV <sup>-2</sup>	0.608	0.011	$b_{-2}$	GeV <sup>-2</sup>	2.935	0.037	$b_{+2}$	GeV <sup>-2</sup>	0.457	0.005	$b_{-2}$	GeV <sup>-2</sup>	2.935	0.014
$b_{+3}$	GeV <sup>-2</sup>	6.45	0.10	$b_{-3}$	GeV <sup>-2</sup>	2.502	0.028	$b_{+3}$	GeV <sup>-2</sup>	7.63	0.02	$b_{-3}$	GeV <sup>-2</sup>	2.502	0.004
$b_P$	GeV <sup>-2</sup>	0.0801	0.0291	$b_O$	GeV <sup>-2</sup>	14.75	0.31	$b_P$	GeV <sup>-2</sup>	$b_P \rightarrow 0$	0	$b_O$	GeV <sup>-2</sup>	14.75	0.05
$b_{PP}$	GeV <sup>-2</sup>	5.287	0.034	$b_{OP}$	GeV <sup>-2</sup>	2.480	0.030	$b_{PP}$	GeV <sup>-2</sup>	4.808	0.007	$b_{OP}$	GeV <sup>-2</sup>	2.480	0.007
$b_R^+$	GeV <sup>-2</sup>	1.928	0.062	$b_R^-$	GeV <sup>-2</sup>	9.25	0.22	$b_R^+$	GeV <sup>-2</sup>	1.596	0.012	$b_R^-$	GeV <sup>-2</sup>	9.25	0.07
$b_{RP}^+$	GeV <sup>-2</sup>	0.4525	0.0215	$b_{RP}^-$	GeV <sup>-2</sup>	1.154	0.039	$b_{RP}^+$	GeV <sup>-2</sup>	0.3862	0.0050	$b_{RP}^-$	GeV <sup>-2</sup>	1.154	0.005
$N_+$	mb GeV <sup>2</sup>	-0.0773	0.0043	$N_-$	mb GeV <sup>2</sup>	15.93	1.54	$N_+$	mb GeV <sup>2</sup>	-0.0508	0.0014	$N_-$	mb GeV <sup>2</sup>	15.93	0.84
$t_+$	GeV <sup>2</sup>	1.475	0.021	$t_-$	GeV <sup>2</sup>	0.1221	0.0088	$t_+$	GeV <sup>2</sup>	1.700	0.013	$t_-$	GeV <sup>2</sup>	0.1221	0.0011
<b>FQ<sub>INT</sub> = 1.598; FQ<sub>EXT</sub> = 1.604; with oderon</b>				$A_{OP}$	GeV <sup>-2</sup>	<b>0.0</b>	(fix)	<b>FQ<sub>INT</sub> = 1.648; FQ<sub>EXT</sub> = 1.694; without oderon (at <math>t=0</math>)</b>				$A_{OP}$	GeV <sup>-2</sup>	<b>0.0</b>	(fix)
				$A_O$	GeV <sup>-2</sup>	-34.72	2.01					$A_O$	GeV <sup>-2</sup>	-34.72	0.52

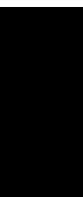
FQ<sub>INT</sub> calculated with “internal” parameter values of machine precision (16 digits) and FQ<sub>EXT</sub> calculated with rounded parameter values as displayed in the table in accordance with PDG rules.

In summary, the solution obtained gives satisfactory picture of the used parametric description of the current database on observables related to elastic (anti)proton–proton scattering amplitudes, and reveals problems with lack of good data at the pre-asymptotic energies. Indeed:

1. Noisy data in dip/shoulder regions does not allow to tune parameters to give credible description of the depth of dips;
2. All frames in Figure 51.12 with  $\sqrt{s} \leq 12$  GeV apparently show that there is an urgent need in  $\bar{p}p$  data at CNI as well as at the first dip/shoulder “ $-t$ ” intervals;
3. Frame marked as “9 to 10 GeV” shows some contradictory data samples in  $pp$  collisions (48 “2.4-std outliers” out of 565 data points). There are no model independent resolution of these contradictions other than remeasurements with much higher statistics and more precise measuring systems. New accurate experimental data are highly desirable;
4. There is a sharp difference in descriptions of  $\rho$  parameter data from our global fits with HPR<sub>1</sub>R<sub>2</sub> model without oderon contribution (non-intersecting  $\rho_{\mp}$  curves) and our variant of AGNM model [7] with oderons (intersecting  $\rho_{\mp}$  curves). Further modelling is needed to remove this difference. Another issue to note is that the  $\rho_-(\sqrt{s})$  curve is consistently above the  $\rho_+(\sqrt{s})$  curve with a constant offset (see Figure 19 in [12]);
5. The 277 ( $> 2.4\sigma$ ) outliers contribute to about one third of the total  $\chi^2$  for the 5009 experimental data points. This issue is not fully understood but may indicate that the procedure of removing some “uncomfortable data” [11,13-15] should be examined more carefully.

## References

- [1] V.M. Abazov *et al.*, Phys. Rev. **D86**, 012009 (2012);
- [2] G. Antchev *et al.*, Nucl. Phys. B **899**, 527 (2015); Phys. Rev. Lett. **111**, 012001 (2013); Europhys. Lett.: **101**, 21004 (2013); **101**, 21002 (2013); **96**, 21002 (2011); **95**, 41001 (2011).
- [3] G. Aad *et al.*, Nucl. Phys. B **889**, 486 (2014).
- [4] P. Abreu *et al.*, Phys. Rev. Lett. **109**, 062002 (2012).
- [5] R. Avila, P. Gauron, and B. Nicolescu, Eur. Phys. J. **C49**, 581 (2007).
- [6] E. Martynov, Phys. Rev. **D76**, 074030 (2007).
- [7] E. Martynov and B. Nicolescu, Eur. Phys. J. **C56**, 57 (2008).
- [8] In Landolt-Börnstein, Group I: P.J. Carson, v.9, 675 (1980); R.R. Shubert, v.9, 216 (1980); P.J. Carlson, v.7, 109 (1973); A.N. Diddens, v.7, 27 (1973).
- [9] J. R. Cudell, A. Lengyel, and E. Martynov, Phys. Rev. **D73**, 034008 (2006).
- [10] R. Cahn, Z. Phys. **C15**, 253 (1982).
- [11] E. Martynov, Phys. Rev. **D87**, 114018 (2013).
- [12] O.V. Selyugin, Phys. Rev. **D91**, 113003 (2015).
- [13] A. Donnachie and P.V. Landshoff, Phys. Lett. **B727**, 500 (2013).
- [14] A.A. Godizov, Phys. Lett. **B735**, 57 (2014).
- [15] A. Donnachie and P.V. Landshoff, Int. J. Mod. Phys. **A29**, 1446007 (2014).



## INTRODUCTION TO THE PARTICLE LISTINGS

Illustrative key . . . . .	601
Abbreviations . . . . .	602







# Illustrative Key to the Particle Listings

Name of particle. "Old" name used before 1986 renaming scheme also given if different. See the section "Naming Scheme for Hadrons" for details.

$a_0(1200)$

$I^G(J^{PC}) = 1^-(0^{++})$

Particle quantum numbers (where known).

OMITTED FROM SUMMARY TABLE  
Evidence not compelling, may be a kinematic effect.

Indicates particle omitted from Particle Physics Summary Table, implying particle's existence is not confirmed.

Quantity tabulated below.

## $a_0(1200)$ MASS

Top line gives our best value (and error) of quantity tabulated here, based on weighted average of measurements used. Could also be from fit, best limit, estimate, or other evaluation. See next page for details.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1206 ± 7 OUR AVERAGE</b>					
1210 ± 8 ± 9	3000	FENNER 87	MMS	-	3.5 $\pi^- p$
1198 ± 10		PIERCE 83	ASPK	+	2.1 $K^- p$
1216 ± 11 ± 9	1500	MERRILL 81	HBC	0	3.2 $K^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1192 ± 16	200	LYNCH 81	HBC	±	2.7 $\pi^- p$

General comments on particle.

"Document id" for this result; full reference given below.

Measurement technique. (See abbreviations on next page.)

Footnote number linking measurement to text of footnote.

<sup>1</sup>Systematic error was added quadratically by us in our 1986 edition.

## $a_0(1200)$ WIDTH

Number of events above background.

Measured value used in averages, fits, limits, etc.

Error in measured value (often statistical only; followed by systematic if separately known; the two are combined in quadrature for averaging and fitting.)

Measured value *not used* in averages, fits, limits, etc. See the Introductory Text for explanations.

Arrow points to weighted average.

Shaded pattern extends  $\pm 1\sigma$  (scaled by "scale factor" S) from weighted average.

Value and error for each experiment.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>41 ± 11 OUR AVERAGE</b>					Error includes scale factor of 1.8. See the ideogram below.
50 ± 8		PIERCE 83	ASPK	+	2.1 $K^- p$
70 + 30 - 20	200	LYNCH 81	HBC	±	2.7 $\pi^- p$
25 ± 5 ± 7		MERRILL 81	HBC	0	3.2 $K^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<60		FENNER 87	MMS	-	3.5 $\pi^- p$

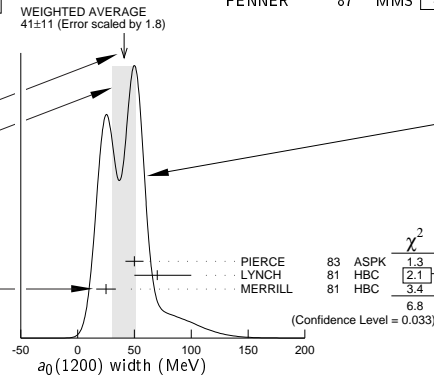
Scale factor > 1 indicates possibly inconsistent data.

Reaction producing particle, or general comments.

"Change bar" indicates result added or changed since previous edition.

Charge(s) of particle(s) detected.

Ideogram to display possibly inconsistent data. Curve is sum of Gaussians, one for each experiment (area of Gaussian = 1/error; width of Gaussian =  $\pm$ error). See Introductory Text for discussion.



Contribution of experiment to  $\chi^2$  (if no entry present, experiment not used in calculating  $\chi^2$  or scale factor because of very large error).

## $a_0(1200)$ DECAY MODES

Partial decay mode (labeled by  $\Gamma_i$ ).

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $3\pi$	(65.2 ± 1.3) %	S=1.7
$\Gamma_2$ $K\bar{K}$	(34.8 ± 1.3) %	S=1.7
$\Gamma_3$ $\eta\pi^\pm$	< 4.9 × 10 <sup>-4</sup>	CL=95%

Our best value for branching fraction as determined from data averaging, fitting, evaluating, limit selection, etc. This list is basically a compact summary of results in the Branching Ratio section below.

## $a_0(1200)$ BRANCHING RATIOS

Branching ratio.

Our best value (and error) of quantity tabulated, as determined from constrained fit (using *all significant* measured branching ratios for this particle).

Weighted average of measurements of this ratio only.

Footnote (referring to LYNCH 81).

$\Gamma(3\pi)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
	<b>0.652 ± 0.013 OUR FIT</b>				Error includes scale factor of 1.7.
	<b>0.643 ± 0.010 OUR AVERAGE</b>				
	0.64 ± 0.01	PIERCE 83	ASPK	+	2.1 $K^- p$
	0.74 ± 0.06	MERRILL 81	HBC	0	3.2 $K^- p$
	• • • We do not use the following data for averages, fits, limits, etc. • • •				
	0.48 ± 0.15	<sup>2</sup> LYNCH 81	HBC	±	2.7 $\pi^- p$
	<sup>2</sup> Data has questionable background subtraction.				

$\Gamma_1/\Gamma$

$\Gamma(K\bar{K})/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
	<b>0.348 ± 0.013 OUR FIT</b>				Error includes scale factor of 1.7.
	<b>0.35 ± 0.05</b>	PIERCE 83	ASPK	+	2.1 $K^- p$

$\Gamma_2/\Gamma$

$\Gamma(K\bar{K})/\Gamma(3\pi)$	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
	<b>0.535 ± 0.030 OUR FIT</b>				Error includes scale factor of 1.7.
	<b>0.50 ± 0.03</b>	MERRILL 81	HBC	0	3.2 $K^- p$

$\Gamma_2/\Gamma_1$

Confidence level for measured upper limit.

$\Gamma(\eta(\text{neutral decay})\pi^\pm)/\Gamma_{\text{total}}$	VALUE (units 10 <sup>-4</sup> )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
	<b>&lt;3.5</b>	95	PIERCE 83	ASPK	+	2.1 $K^- p$

0.71  $\Gamma_3/\Gamma$

References, ordered inversely by year, then author.

"Document id" used on data entries above.

Journal, report, preprint, etc. (See abbreviations on next page.)

## $a_0(1200)$ REFERENCES

FENNER 87	PRL 55 14	H. Fenner et al.	(SLAC)
PIERCE 83	PL 123B 230	J.H. Pierce	(FNAL) UP
LYNCH 81	PR D24 610	G.R. Lynch et al.	(CLEO Collab.)
MERRILL 81	PRL 47 143	D.W. Merrill et al.	(SACL, CERN)

Partial list of author(s) in addition to first author.

Quantum number determinations in this reference.

Institution(s) of author(s). (See abbreviations on next page.)

# Abbreviations Used in the Particle Listings

## Indicator of Procedure Used to Obtain Our Result

OUR AVERAGE	From a weighted average of selected data.
OUR FIT	From a constrained or overdetermined multiparameter fit of selected data.
OUR EVALUATION	Not from a direct measurement, but evaluated from measurements of other quantities.
OUR ESTIMATE	Based on the observed range of the data. Not from a formal statistical procedure.
OUR LIMIT	For special cases where the limit is evaluated by us from measured ratios or other data. Not from a direct measurement.

## Measurement Techniques

(i.e., Detectors and Methods of Analysis)

A1	A1 Collaboration at MAMI
ACCM	ACCMOR Collaboration
ADMX	Axion Dark Matter Experiment
AEMS	Argonne effective mass spectrometer
ALEP	ALEPH – CERN LEP detector
ALPS	Photon regeneration experiment
AMND	AMANDA South Pole neutrino detector
AMY	AMY detector at KEK-TRISTAN
ANIT	Antarctic Impulsive Transient Antenna balloon mission
ANTR	ANTARES underwater neutrino telescope in the Western Mediterranean Sea
APEX	FNAL APEX Collab.
ARG	ARGUS detector at DORIS
ARGD	Fit to semicircular amplitude path on Argand diagram
ASP	Anomalous single-photon detector
ASPK	Automatic spark chambers
ASTE	ASTERIX detector at LEAR
ASTR	Astronomy
ATLS	ATLAS detector at CERN LHC
B787	BNL experiment 787 detector
B791	BNL experiment 791 detector
B845	BNL experiment 845 detector
B852	BNL E-852
B865	BNL E865 detector
B871	BNL experiment 871 detector
B949	BNL E949 detector at AGS
BABR	BaBar Collab.
BAIK	Lake Baikal neutrino telescope
BAKS	Baksan underground scintillation telescope
BC	Bubble chamber
BDMP	Beam dump
BEAT	CERN BEATRICE Collab.
BEBC	Big European bubble chamber at CERN
BELL	Belle Collab.
BES	BES Beijing Spectrometer at Beijing Electron-Positron Collider
BES2	BES Beijing Spectrometer at Beijing Electron-Positron Collider
BES3	BES Beijing Spectrometer at Beijing Electron-Positron Collider
BIS2	BIS-2 spectrometer at Serpukhov
BKEI	BENKEI spectrometer system at KEK Proton Synchrotron
BOLO	Bolometer, a cryogenic thermal detector
BONA	Bonanza nonmagnetic detector at DORIS
BORX	BOREXINO
BPWA	Barrelet-zero partial-wave analysis
CALO	Calorimeter
CAST	CAST experiment at CERN
CBAL	Crystal Ball detector at SLAC-SPEAR or DORIS
CBAR	Crystal Barrel detector at CERN-LEAR
CBOX	Crystal Box at LAMPF
CBTP	CBELSA/TAPS Collaboration
CC	Cloud chamber
CCFR	Columbia-Chicago-Fermilab-Rochester detector
CDEX	China Dark Matter Experiment
CDF	Collider detector at Fermilab
CDF2	CDF-II Collab.
CDHS	CDHS neutrino detector at CERN
CDM2	CDMS II, Cryogenic Dark Matter Search at Soudan Underground Lab.
CDMS	CDMS Collaboration
CELL	CELLO detector at DESY
CGNT	CoGeNT dark matter search experiment
CHER	Cherenkov detector
CHM2	CHARM-II neutrino detector (glass) at CERN
CHOZ	Nuclear Power Station near Chooz, France
CHRM	CHARM neutrino detector (marble) at CERN

CHRS	CHORUS Collaboration – CERN SPS
CIB	Cosmic Infrared Background
CIBS	CERN-IHEP boson spectrometer
CLAS	Jefferson CLAS Collab.
CLE2	CLEO II detector at CESR
CLE3	CLEO III detector at CESR
CLEC	CLEO-c detector at CESR
CLEO	Cornell magnetic detector at CESR
CMB	Cosmic Microwave Background
CMD	Cryogenic magnetic detector at VEPP-2M, Novosibirsk
CMD2	Cryogenic magnetic detector 2 at VEPP-2M, Novosibirsk
CMD3	Cryogenic magnetic detector 3 at VEPP-2000, Novosibirsk
CMS	CMS detector at CERN LHC
CNTR	Counters
COMB	Combined analysis of data from independent experiments.
COMP	COMPASS experiment at the CERN SPS
COSM	Cosmology and astrophysics
COSY	COSY-TOF Collaboration
COUP	COUPP (the Chicagoland Observatory for Underground Particle Physics) Collab.
CPLR	CPLEAR Collaboration
CRBT	Crystal Ball and TAPS detector at MAMI
CRES	CRESST cryogenic detector
CRYB	Crystal Ball at BNL
CRYM	Crystal Ball detector at Mainz Microtron MAMI
CSB2	Columbia U. - Stony Brook BGO calorimeter inserted in NaI array
CSME	COSME Collaboration
CUOR	CUORICINO experiment at Gran Sasso Laboratory.
CUSB	Columbia U. - Stony Brook segmented NaI detector at CESR
D0	D0 detector at Fermilab Tevatron Collider
DAMA	DAMA, dark matter detector at Gran Sasso National Lab.
DASP	DESY double-arm spectrometer
DAYA	Daya Bay Collaboration
DBC	Deuterium bubble chamber
DCHZ	Double Chooz Collaboration
DLCO	DELCO detector at SLAC-SPEAR or SLAC-PEP
DLPH	DELPHI detector at LEP
DM1	Magnetic detector no. 1 at Orsay DCI collider
DM2	Magnetic detector no. 2 at Orsay DCI collider
DMIC	DAMIC Dark Matter in CCD experiment at Fermilab
DMTP	Dark Matter Time Projection Chamber (DMTPC) directional detection experiment
DONU	DONUT Collab.
DPWA	Energy-dependent partial-wave analysis
DRFT	Directional dark matter detector at Boulby Underground Science Facility
DS50	DarkSide-50 Liquid Argon TPC at Gran Sasso National Laboratory
E621	Fermilab E621 detector
E653	Fermilab E653 detector
E665	Fermilab E665 detector
E687	Fermilab E687 detector
E691	Fermilab E691 detector
E705	Fermilab E705 Spectrometer-Calorimeter
E731	Fermilab E731 Spectrometer-Calorimeter
E756	Fermilab E756 detector
E760	Fermilab E760 detector
E761	Fermilab E761 detector
E771	Fermilab E771 detector
E773	Fermilab E773 Spectrometer-Calorimeter
E789	Fermilab E789 detector
E791	Fermilab E791 detector
E799	Fermilab E799 Spectrometer-Calorimeter
E835	Fermilab E835 detector
EDE2	EDELWEISS II dark matter search Collaboration
EDEL	EDELWEISS dark matter search Collaboration
EHS	Four-pi detector at CERN
ELEC	Electronic combination
EMC	European muon collaboration detector at CERN
EMUL	Emulsions
FAST	Fiber Active Scintillator Target detector at PSI
FBC	Freon bubble chamber
FENI	FENICE (at the ADONE collider of Frascati)
FIT	Fit to previously existing data
FLAT	Large Area Telescope onboard the Fermi Gamma-Ray Space Telescope (Fermi-LAT)
FMPS	Fermilab Multiparticle Spectrometer
FOCS	FNAL E831 FOCUS Collab.
FRAB	ADONE $B\bar{B}$ group detector

# Abbreviations Used in the Particle Listings

FRAG	ADONE $\gamma\gamma$ group detector	MAC	MAC detector at PEP/SLAC
FRAM	ADONE MEA group detector	MBOO	Fermilab MiniBooNE neutrino experiment
FREJ	FREJUS Collaboration – modular flash chamber detector (calorimeter)	MBR	Molecular beam resonance technique
GA24	Hodoscope Cherenkov $\gamma$ calorimeter (IHEP GAMS-2000) (CERN GAMS-4000)	MCRO	MACRO detector in Gran Sasso
GALX	GALLEX solar neutrino detector in the Gran Sasso Underground Lab.	MD1	Magnetic detector at VEPP-4, Novosibirsk
GAM2	IHEP hodoscope Cherenkov $\gamma$ calorimeter GAMS-2000	MDRP	Millikan drop measurement
GAM4	CERN hodoscope Cherenkov $\gamma$ calorimeter GAMS-4000	MEG	Muon to electron conversion detector at PSI
GAMS	IHEP hodoscope Cherenkov $\gamma$ calorimeter GAMS-4 $\pi$	MGIC	MAGIC Telescopes gamma-ray observatory.
GNO	Gallium Neutrino Observatory in the Gran Sasso Underground Lab.	MICA	Underground mica deposits
GOLI	CERN Goliath spectrometer	MINS	Fermilab MINOS experiment
GRAL	GRAAL Collaboration	MIRA	MIRABELLE Liquid-hydrogen bubble chamber
H1	H1 detector at DESY/HERA	MLEV	Magnetic levitation
HBC	Hydrogen bubble chamber	MLS	Modified Laurent Series
HDBC	Hydrogen and deuterium bubble chambers	MMS	Missing mass spectrometer
HDES	HADES Collaboration at GSI in Darmstadt	MPS	Multiparticle spectrometer at BNL
HDMO	Heidelberg-Moscow Experiment	MPS2	Multiparticle spectrometer upgrade at BNL
HDMS	Heidelberg Dark Matter Search Experiment	MPSF	Multiparticle spectrometer at Fermilab
HEBC	Helium bubble chamber	MPWA	Model-dependent partial-wave analysis
HEPT	Helium proportional tubes	MRK1	SLAC Mark-I detector
HERB	HERA-B detector at DESY/HERA	MRK2	SLAC Mark-II detector
HERM	HERMES detector at DESY/HERA	MRK3	SLAC Mark-III detector
HESS	High Energy Stereoscopic System gamma-ray instrument	MRKJ	Mark-J detector at DESY
HFS	Hyperfine structure	MRS	Magnetic resonance spectrometer
HLBC	Heavy-liquid bubble chamber	MUG2	MUON(g-2)
HOME	Homestake underground scintillation detector	MWPC	Multi-Wire Proportional Chamber
HPGE	High-purity Germanium detector	NA14	CERN NA14
HPW	Harvard-Pennsylvania-Wisconsin detector	NA31	CERN NA31 Spectrometer-Calorimeter
HRS	SLAC high-resolution spectrometer	NA32	CERN NA32 Spectrometer
HYBR	Hybrid: bubble chamber + electronics	NA48	CERN NA48 Collaboration
HYCP	HyperCP Collab. (FNAL E-871)	NA49	CERN NA49 Collaboration
IACT	Imaging Air Cherenkov Telescope	NA60	CERN NA60 Collaboration
ICAR	ICARUS experiment at Gran Sasso Laboratory.	NA62	CERN NA62 Experiment
ICCB	IceCube neutrino detector at South Pole	NAGE	NEWAGE, New generation WIMP-search experiment with advanced gaseous tracking
IGEX	IGEX Collab.	NAIA	NAIAD (NaI Advanced Detector) dark matter search experiment
IMB	Irvine-Michigan-Brookhaven underground Cherenkov detector	ND	NaI detector at VEPP-2M, Novosibirsk
IMB3	Irvine-Michigan-Brookhaven underground Cherenkov detector	NICE	Serpukhov nonmagnetic precision spectrometer
INDU	Magnetic induction	NMR	Nuclear magnetic resonance
IPWA	Energy-independent partial-wave analysis	NOMD	NOMAD Collaboration, CERN SPS
ISTR	IHEP ISTRA+ spectrometer-calorimeter	NTEV	NuTeV Collab. at Fermilab
JADE	JADE detector at DESY	nTRV	neutron Time-Reversal Violation
K246	KEK E246 detector with polarimeter	NUSX	Mont Blanc NUSEX underground detector
K2K	KEK to Super-Kamiokande	OBLX	OBELIX detector at LEAR
K391	KEK E391a detector	OLYA	Detector at VEPP-2M and VEPP-4, Novosibirsk
K470	KEK-E470 Stopping K detector	OMEG	CERN OMEGA spectrometer
KAM2	KAMIOKANDE-II underground Cherenkov detector	OPAL	OPAL detector at LEP
KAMI	KAMIOKANDE underground Cherenkov detector	OPER	OPERA experiment with emulsion tracking at Gran Sasso
KAR2	KARMEN2 calorimeter at the ISIS neutron spallation source at Rutherford	OSPK	Optical spark chamber
KARM	KARMEN calorimeter at the ISIS neutron spallation source at Rutherford	PANX	PandaX dual-phase liquid xenon dark matter experiment at Jin-Ping
KEDR	detector operating at VEPP-4M collider (Novosibirsk)	PIBE	The PIBETA detector at the Paul Scherrer Institute (PSI), Switzerland.
KIMS	Korea Invisible Mass Search experiment at YangYang, Korea	PICA	PICASSO dark matter search experiment
KLND	KamLand Collab. (Japan)	PICO	PICO bubble chamber experiment in SNOLAB underground laboratory
KLOE	KLOE detector at DAFNE (the Frascati e <sup>+</sup> e <sup>-</sup> collider Italy)	PIE3	$\pi$ E3 beam-line of Paul Scherrer Institute
KOLR	Kolar Gold Field underground detector	PLAS	Plastic detector
KTEV	KTeV Collaboration	PLUT	DESY PLUTO detector
L3	L3 detector at LEP	PMLA	PAMELA space spectrometer on Resurs-DK1 satellite
LASR	Laser	PRMX	The PRIMEX detector in Hall B at TJNAF
LASS	Large-angle superconducting solenoid spectrometer at SLAC	PWA	Partial-wave analysis
LATT	Lattice calculations	REDE	Resonance depolarization
LEBC	Little European bubble chamber at CERN	RENO	RENO Collaboration
LEGS	BNL LEGS Collab.	RICE	Radio Ice Cherenkov Experiment
LENA	Nonmagnetic lead-glass NaI detector at DORIS	RVUE	Review of previous data
LEP	From combination of all 4 LEP experiments: ALEPH, DELPHI, L3, OPAL	SAGE	US - Russian Gallium Experiment
LEPS	Low-Energy Pion Spectrometer at the Paul Scherrer Institute	SCDM	SuperCDMS experiment at Soudan Underground Lab.
LGW	Lead Glass Wall collaboration at SPEAR/SLAC	SELX	FNAL SELEX Collab.
LHC	Combined analysis of LHC experiments	SFM	CERN split-field magnet
LHCB	LHCb detector at CERN LHC	SHF	SLAC Hybrid Facility Photon Collaboration
LSD	Mont Blanc liquid scintillator detector	SIGM	Serpukhov CERN-IHEP magnetic spectrometer (SIGMA)
LSND	Liquid Scintillator Neutrino Detector	SILI	Silicon detector
LSW	Light Shining through a Wall	SIMP	SIMPLE, dark matter detector at Laboratori Nazionali del Sud
LUX	Large Underground Xenon experiment at SURF	SKAM	Super-Kamiokande Collab.
		SLAX	Solar Axion Experiment in Canfranc Underground Laboratory
		SLD	SLC Large Detector for e <sup>+</sup> e <sup>-</sup> colliding beams at SLAC
		SMPL	SIMPLE, Superheated Instrument for Massive Particle Experiments

## Abbreviations Used in the Particle Listings

SND	Novosibirisk Spherical neutral detector at VEPP-2M	BAPS	Bulletin of the American Physical Society
SNDR	SINDRUM spectrometer at PSI	BASUP	Bulletin of the Academy of Science, USSR (Physics)
SNO	SNO Collaboration (Sudbury Neutrino Observatory)	CJNP	Chinese Journal of Nuclear Physics
SOU2	Soudan 2 underground detector	CJP	Canadian Journal of Physics
SOUT	Soudan underground detector	CNPP	Comments on Nuclear and Particle Physics
SPEC	Spectrometer	CP	Chinese Physics
SPED	From maximum of speed plot or resonant amplitude	CPC	Chinese Physics C
SPHR	Bonn SAPHIR Collab.	CTP	Communications in Theoretical Physics
SPNX	SPHINX spectrometer at IHEP accelerator	CZJP	Czechoslovak Journal of Physics
SPRK	Spark chamber	DANS	Doklady Akademii nauk SSSR
SQID	SQUID device	EPJ	The European Physical Journal
STRC	Streamer chamber	EPL	Europhysics Letters
SVD2	SVD-2 experiment at IHEP, Protvino	FECAF	Fizika Elementarnykh Chastits i Atomnogo Yadra
T2K	T2K Collaboration	HADJ	Hadronic Journal
TASS	DESY TASSO detector	IJMP	International Journal of Modern Physics
TEVA	Combined analysis of CDF and DØ experiments	JAP	Journal of Applied Physics
TEXO	TEXONO Collab., ultra low energy Ge detector at Kuo-Sheng Laboratory	JCAP	Journal of Cosmology and Astroparticle Physics
THEO	Theoretical or heavily model-dependent result	JETP	English Translation of Soviet Physics ZETF
TNF	TNF-IHEP facility at 70 GeV IHEP accelerator	JETPL	English Translation of Soviet Physics ZETF Letters
TOF	Time-of-flight	JHEP	Journal of High Energy Physics
TOPZ	TOPAZ detector at KEK-TRISTAN	JINR	Joint Inst. for Nuclear Research
TPC	TPC detector at PEP/SLAC	JINRRC	JINR Rapid Communications
TPS	Tagged photon spectrometer at Fermilab	JP	Journal of Physics
TRAP	Penning trap	JPA	Journal of Physics, A
TWST	TWIST spectrometer at TRIUMF	JPB	Journal of Physics, B
UA1	UA1 detector at CERN	JPCRD	Journal of Physical and Chemical Reference Data
UA2	UA2 detector at CERN	JPCS	Journal of Physics: Conference Series
UA5	UA5 detector at CERN	JPG	Journal of Physics, G
UCNA	UCNA collaboration using polarized ultracold neutrons at LANSCE	JPSJ	Journal of the Physical Society of Japan
UKDM	UK Dark Matter Collab.	LNC	Lettère Nuovo Cimento
VES	Vertex Spectrometer Facility at 70 GeV IHEP accelerator	MNRAS	Monthly Notices of the Royal Astronomical Society
VLBI	Very Long Baseline Interferometer	MPL	Modern Physics Letters
VNS	VENUS detector at KEK-TRISTAN	NAST	New Astronomy
VRTS	Very Energetic Radiation Imaging Telescope Array System (VERITAS)	NAT	Nature
WA75	CERN WA75 experiment	NATP	Nature Physics
WA82	CERN WA82 experiment	NC	Nuovo Cimento
WA89	CERN WA89 experiment	NIM	Nuclear Instruments and Methods
WARP	Liquid argon detector for CDM searches at Gran Sasso	NJP	New Journal of Physics
WASA	WASA detector at CELSIUS, Uppsala and at COSY, Juelich	NP	Nuclear Physics
WIRE	Wire chamber	NPBPS	Nuclear Physics B Proceedings Supplement
X100	XENON100 dark matter search experiment at Gran Sasso National Laboratory	PAN	Physics of Atomic Nuclei (formerly SJNP)
XE10	XENON10 experiment at Gran Sasso National Laboratory	PD	Physics Doklady (Magazine)
XEBC	Xenon bubble chamber	PDAT	Physik Daten
XMAS	XMASS, liquid xenon scintillation detector at Kamioka Observatory	PL	Physics Letters
ZEP2	ZEPLIN-II dark matter detector	PN	Particles and Nuclei
ZEP3	ZEPLIN-III dark matter detector at Palmer Underground Lab.	PPCF	Plasma Physics and Controlled Fusion
ZEPL	ZEPLIN-I galactic dark matter detector	PPN	Physics of Particles and Nuclei (formerly SJNP)
ZEUS	ZEUS detector at DESY/HERA	PPNL	Physics of Particles and Nuclei Letters
		PPNP	Progress in Particles and Nuclear Physics
		PPSL	Proc. of the Physical Society of London
		PR	Physical Review
		PRAM	Pramana
		PRL	Physical Review Letters
		PRPL	Physics Reports (Physics Letters C)
		PRSE	Proc. of the Royal Society of Edinburgh
		PRSL	Proc. of the Royal Society of London, Section A
		PS	Physica Scripta
		PTEP	Progress of Theoretical and Experimental Physics
		PTP	Progress of Theoretical Physics
		PTPS	Progress of Theoretical Physics Supplement
		PTRSL	Phil. Trans. Royal Society of London
		RA	Radiochimica Acta
		RMP	Reviews of Modern Physics
		RNC	La Rivista del Nuovo Cimento
		RPP	Reports on Progress in Physics
		RRP	Revue Roumaine de Physique
		SCI	Science
		SJNP	Soviet Journal of Nuclear Physics
		SJPN	Soviet Journal of Particles and Nuclei
		SPD	Soviet Physics Doklady (Magazine)
		SPU	Soviet Physics - Uspekhi
		UFN	Usp. Fiz. Nauk – Russian version of SPU
		YAF	Yadernaya Fizika
		ZETF	Zhurnal Eksperimental'noi i Teoreticheskoi Fiziki
		ZETFP	Zhurnal Eksperimental'noi i Teoreticheskoi Fiziki, Pis'ma v Redakts

### Conferences

Conferences are generally referred to by the location at which they were held (e.g., HAMBURG, TORONTO, CORNELL, BRIGHTON, etc.).

### Journals

AA	Astronomy and Astrophysics
ADVP	Advances in Physics
AFIS	Anales de Fisica
AJP	American Journal of Physics
AL	Astronomy Letters
ANP	Annals of Physics
ANPL	Annals of Physics (Leipzig)
ANYAS	Annals of the New York Academy of Sciences
AP	Atomic Physics
APAH	Acta Physica Academiae Scientiarum Hungaricae
APJ	Astrophysical Journal
APJS	Astrophysical Journal Suppl.
APP	Acta Physica Polonica
APS	Acta Physica Slovaca
ARNPS	Annual Review of Nuclear and Particle Science
ARNS	Annual Review of Nuclear Science
ASP	Astroparticle Physics
AST	American Statistician

# Abbreviations Used in the Particle Listings

ZNAT	Zeitschrift für Naturforschung		BERL	<b>DESY</b> , Deutsches Elektronen-Synchrotron	<b>Zeuthen</b> , Germany
ZPHY	Zeitschrift für Physik		BERN	Univ. of <b>Berne</b>	Berne, Switzerland
<b>Institutions</b>			BGNA	Univ. di <b>Bologna</b> , & <b>INFN</b> , Sezione di Bologna; Via Irnerio, 46, I-40126 Bologna; Viale C. Berti Pichat, n. 6/2	Bologna, Italy
AACH	Phys. Inst. der Techn. Hochschule <b>Aachen</b> (Historical, use for general Inst. der Techn. Hochschule)	Aachen, Germany	BHAB	<b>Bhabha Atomic Research Center</b>	Trombay, Bombay, India
AACH1	I Phys. Inst. B, RWTH <b>Aachen</b>	Aachen, Germany	BHEP	<b>Inst. of High Energy Physics</b>	<b>Beijing</b> , China
AACH3	III Phys. Inst. A, RWTH <b>Aachen</b> Univ.	Aachen, Germany	BIEL	Univ. <b>Bielefeld</b>	Bielefeld, Germany
AACHT	Inst. für Theoretische Teilchenphysik & Kosmologie, RWTH <b>Aachen</b>	Aachen, Germany	BING	<b>SUNY at Binghamton</b>	Binghamton, NY, USA
AARH	Univ. of <b>Aarhus</b>	Aarhus C, Denmark	BIRK	<b>Birkbeck College</b> , Univ. of London	London, United Kingdom
ABO	<b>Åbo</b> Akademi Univ.	Turku, Finland	BIRM	Univ. of <b>Birmingham</b>	Edgbaston, Birmingham, United Kingdom
ADEL	<b>Adelphi</b> Univ.	Garden City, NY, USA	BLSU	<b>Bloomsburg</b> Univ.	Bloomsburg, PA, USA
ADLD	The Univ. of <b>Adelaide</b>	Adelaide, SA, Australia	BNL	<b>Brookhaven National Lab.</b>	Upton, NY, USA
AERE	Atomic Energy Research Estab.	Didcot, United Kingdom	BOCH	<b>Ruhr</b> Univ. <b>Bochum</b>	Bochum, Germany
AFRR	<b>Armed Forces</b> Radiobiology Res. Inst.	Bethesda, MD, USA	BOHR	<b>Niels Bohr Inst.</b>	Copenhagen Ø, Denmark
AHMED	Physical Research Lab.	<b>Ahmedabad</b> , Gujarat, India	BOIS	<b>Boise</b> State Univ.	Boise, ID, USA
AICH	<b>Aichi</b> Univ. of Education	Aichi, Japan	BOMB	Univ. of <b>Bombay</b>	Bombay, India
AKIT	<b>Akita</b> Univ.	Akita, Japan	BONN	<b>Univ. of Bonn</b>	Bonn, Germany
ALAH	Univ. of <b>Alabama</b> ( <b>Huntsville</b> )	Huntsville, AL, USA	BORD	Centre d'Etudes Nucléaires de <b>Bordeaux</b> Gradignan ( <b>CENBG</b> )	Gradignan, France
ALAT	Univ. of <b>Alabama</b> ( <b>Tuscaloosa</b> )	Tuscaloosa, AL, USA	BOSE	S.N. <b>Bose</b> National Centre for Basis Sciences	Calcutta, India
ALBA	<b>SUNY at Albany</b>	Albany, NY, USA	BOSK	<b>"Rudjer Bošković"</b> Inst.	Zagreb, Croatia
ALBE	Univ. of <b>Alberta</b>	Edmonton, AB, Canada	BOST	<b>Boston</b> Univ.	Boston, MA, USA
AMES	<b>Ames Lab.</b>	Ames, IA, USA	BRAN	<b>Brandeis</b> Univ.	Waltham, MA, USA
AMHT	<b>Amherst</b> College	Amherst, MA, USA	BRCO	Univ. of <b>British Columbia</b>	Vancouver, BC, Canada
AMST	Univ. van <b>Amsterdam</b>	GL Amsterdam, The Netherlands	BRIS	Univ. of <b>Bristol</b>	Bristol, United Kingdom
ANIK	<b>NIKHEF</b>	<b>Amsterdam</b> , The Netherlands	BROW	<b>Brown</b> Univ.	Providence, RI, USA
ANKA	<b>Middle East</b> Technical Univ.; Dept. of Physics; Experimental HEP Lab	Ankara, Turkey	BRUN	<b>Brunel</b> Univ.	Uxbridge, Middlesex, United Kingdom
ANL	<b>Argonne</b> National Lab.; High Energy Physics Division, Bldg. 362; Physics Division, Bldg. 203	Argonne, IL, USA	BRUX	Univ. Libre de <b>Bruxelles</b> ; Physique des Particules Élémentaires	Bruxelles, Belgium
ANSM	<b>St. Anselm</b> Coll.	Manchester, NH, USA	BRUXT	Univ. Libre de <b>Bruxelles</b> ; Physique Théorique	Bruxelles, Belgium
ARCBO	<b>Arecibo</b> Observatory	Arecibo, PR, USA	BUCH	Univ. of <b>Bucharest</b>	Bucharest-Magurele, Romania
ARIZ	Univ. of <b>Arizona</b>	Tucson, AZ, USA	BUDA	<b>Wigner</b> Research Centre for Physics	<b>Budapest</b> , Hungary
ARZS	<b>Arizona State</b> Univ.	Tempe, AZ, USA	BUFF	<b>SUNY at Buffalo</b>	Buffalo, NY, USA
ASCI	Russian Academy of Sciences	<b>Moscow</b> , Russian Federation	BURE	Inst. des Hautes Etudes Scientifiques	<b>Bures-sur-Yvette</b> , France
AST	<b>Academia Sinica</b>	Nankang, Taipei, Taiwan	CAEN	Lab. de Physique Corpusculaire, <b>ENSICAEN</b>	<b>Caen</b> , France
ATEN	NCSR <b>"Demokritos"</b>	<b>Aghia</b> Paraskevi, Greece	CAGL	Univ. degli Studi di <b>Cagliari</b>	Monsserrato (CA), Italy
ATHU	Univ. of <b>Athens</b>	Athens, Greece	CAIR	<b>Cairo</b> University	Orman, Giza, Cairo, Egypt
AUCK	Univ. of <b>Auckland</b>	Auckland, New Zealand	CAIW	<b>Carnegie Inst.</b> of Washington	Washington, DC, USA
BAKU	<b>Natl. Azerbaijan Academy of Sciences</b> , Inst. of Physics	<b>Baku</b> , Azerbaijan	CALB	Univ. della <b>Calabria</b>	Cosenza, Italy
BANG	Indian Inst. of Science	<b>Bangalore</b> , India	CALC	Univ. of <b>Calcutta</b>	Calcutta, India
BANGB	<b>Bangabasi</b> College	Calcutta, India	CAMB	<b>DAMTP</b>	Cambridge, United Kingdom
BARC	Univ. Autònoma de <b>Barcelona</b>	Bellaterra (Barcelona), Spain	CAMP	Univ. Estadual de <b>Campinas</b> (UNICAMP)	<b>Campinas</b> , SP, Brasil
BARI	Univ. e del Politecnico di <b>Bari</b>	Bari, Italy	CANB	<b>Australian National Univ.</b>	Canberra, ACT, Australia
BART	Univ. of <b>Delaware</b> ; <b>Bartol</b> Research Inst.	Newark, DE, USA	CANTB	Inst. de Física de Cantabria (CSIC-Univ. <b>Cantabria</b> )	Santander, Spain
BASL	Inst. für Physik der Univ. <b>Basel</b>	Basel, Switzerland	CAPE	University of <b>Cape Town</b>	Rondebosch, Cape Town, South Africa
BAYR	Univ. <b>Bayreuth</b>	Bayreuth, Germany	CARA	Univ. Central de <b>Venezuela</b>	Caracas, Venezuela
BCEN	Centre d'Etudes Nucleaires de <b>Bordeaux-Gradignan</b>	Gradignan, France	CARL	<b>Carleton</b> Univ.	Ottawa, ON, Canada
BCIP	Natl. Inst. for Physics & Nuclear Eng. "Horia Hulubei" ( <b>IFIN-HH</b> )	<b>Bucharest</b> -Magurele, Romania	CARLC	<b>Carleton</b> College	Northfield, MN, USA
BEIJ	<b>Beijing</b> Univ.	Beijing, China	CASE	<b>Case</b> Western Reserve Univ.	Cleveland, OH, USA
BEIJT	<b>Inst. of Theoretical Physics</b>	<b>Beijing</b> , China	CAST	<b>China</b> Center of Advanced Science and Technology	Beijing, China
BELG	Inter-University Inst. for High Energies (ULB-VUB)	<b>Brussel</b> , Belgium	CATA	Univ. di <b>Catania</b>	Catania, Italy
BELL	AT & T <b>Bell</b> Labs	Murray Hill, NJ, USA	CATH	<b>Catholic</b> Univ. of America	Washington, DC, USA
BERG	Univ. of <b>Bergen</b>	Bergen, Norway	CAVE	<b>Cavendish</b> Lab.	Cambridge, United Kingdom
			CBNM	CBNM	<b>Geel</b> , Belgium
			CBPF	Centro Brasileiro de Pesquisas Físicas – <b>BIB/CDI/CBPF</b>	<b>Rio de Janeiro</b> , RJ, Brasil
			CCAC	<b>Allegheny</b> College	Meadville, PA, USA
			CDEF	<b>Univ. Paris VII, Denis Diderot</b>	Paris, France

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CEA	Cambridge Electron Accelerator (Historical in <i>Review</i> )	Cambridge, MA, USA	FERR	Univ. di <b>Ferrara</b>	Ferrara, Italy
CEADE	Center for Apl. Studies for Nuclear Physics	Havana, Cuba	FIRZ	Univ. degli Studi di <b>Firenze</b>	Sesto Fiorentino, Italy
CEBAF	<b>Jefferson Lab—Thomas Jefferson National Accelerator Facility</b>	Newport News, VA, USA	FISK	<b>Fisk</b> Univ.	Nashville, TN, USA
CENG	Centre d'Etudes Nucleaires	Grenoble, France	FLOR	Univ. of <b>Florida</b>	Gainesville, FL, USA
CERN	<b>CERN</b> , European Organization for Nuclear Research	Genève, Switzerland	FNAL	<b>Fermilab</b>	Batavia, IL, USA
CFPA	Univ. of <b>California</b> , (Berkeley)	Berkeley, CA, USA	FOM	<b>FOM</b> , Stichting voor Fundamenteel Onderzoek der Materie	JP <b>Utrecht</b> , The Netherlands
CHIC	Univ. of <b>Chicago</b>	Chicago, IL, USA	FRAN	<b>Frankfurt Inst. for Advanced Studies (FIAS)</b>	Frankfurt am Main, Germany
CIAE	<b>State Nuclear Power Research Inst.</b>	Beijing, China	FRAS	Lab. Nazionali di <b>Frascati</b> dell'INFN	Frascati (Roma), Italy
CINC	Univ. of <b>Cincinnati</b>	Cincinnati, OH, USA	FREIB	<b>Albert-Ludwigs</b> Univ.	<b>Freiburg</b> , Germany
CINV	<b>CINVESTAV-IPN</b> Centro de Investigacion y de Estudios Avanzados del IPN	México, DF, Mexico	FREIE	<b>Freie</b> Univ. <b>Berlin</b>	Berlin, Germany
CIT	<b>California Inst. of Tech.</b>	Pasadena, CA, USA	FRIB	Univ. de <b>Fribourg</b>	Fribourg, Switzerland
CLER	Univ. de <b>Clermont-Ferrand</b>	Aubière, France	FSU	<b>Florida State</b> Univ.; High Energy Physics	Tallahassee, FL, USA
CLEV	<b>Cleveland</b> State Univ.	Cleveland, OH, USA	FSUSC	<b>Florida State</b> Univ.; SCS (School of Computational Science)	Tallahassee, FL, USA
CMNS	<b>Comenius</b> Univ. (FMFI UK)	Bratislava, Slovakia	FUKI	<b>Fukui</b> Univ.	Fukui, Japan
CMU	<b>Carnegie Mellon</b> Univ.	Pittsburgh, PA, USA	FUKU	<b>Fukushima</b> Univ.	Fukushima, Japan
CNEA	<b>Comisión Nacional de Energía Atómica</b>	Buenos Aires, Argentina	GENO	Univ. di <b>Genova</b>	Genova, Italy
CNRC	Centre for Research in Particle Physics	Ottawa, ON, Canada	GEOR	<b>E. Andronikashvili Inst. of Physics</b>	Tbilisi, Republic of Georgia
COIM	Univ. de <b>Coimbra</b>	Coimbra, Portugal	GESC	<b>General Electric</b> Co.	Schenectady, NY, USA
COLO	Univ. of <b>Colorado</b>	Boulder, CO, USA	GEVA	Univ. de <b>Genève</b>	Genève, Switzerland
COLU	<b>Columbia</b> Univ.	New York, NY, USA	GIES	Univ. <b>Giessen</b>	Giessen, Germany
CONC	<b>Concordia</b> University	Montreal, PQ, Canada	GIFU	<b>Gifu</b> Univ.	Gifu, Japan
CORN	<b>Cornell</b> Univ.	Ithaca, NY, USA	GLAS	Univ. of <b>Glasgow</b>	Glasgow, United Kingdom
COSU	<b>Colorado State</b> Univ.	Fort Collins, CO, USA	GMAS	<b>George Mason</b> Univ.	Fairfax, VA, USA
CPPM	Centre National de la Recherche Scientifique, Luminy	Marseille, France	GOET	Univ. <b>Göttingen</b>	Göttingen, Germany
CRAC	Henryk Niewodnicza'nski Inst. of Nuclear Physics	Kraków, Poland	GRAN	Univ. de <b>Granada</b>	Granada, Spain
CRNL	<b>Chalk River</b> Labs.	Chalk River, ON, Canada	GRAZ	Univ. <b>Graz</b>	Graz, Austria
CSOK	<b>Oklahoma</b> Central State Univ.	Edmond, OK, USA	GRON	Univ. of <b>Groningen</b>	Groningen, The Netherlands
CST	Univ. of <b>Science and Technology</b> of China	Hefei, Anhui 230026, China	GSCO	<b>Geological Survey of Canada</b>	Ottawa, ON, Canada
CSULB	<b>California</b> State Univ.	Long Beach, CA, USA	GSI	GSI Helmholtzzentrum für Schwerionenforschung GmbH	<b>Darmstadt</b> , Germany
CSUS	<b>California</b> State Univ.	Sacramento, CA, USA	GUAN	Univ. de <b>Guanajuato</b>	León, Gto., Mexico
CUNY	<b>City College</b> of <b>New York</b>	New York, NY, USA	GUEL	Univ. of <b>Guelph</b>	Guelph, ON, Canada
CURCP	<b>Univ. Pierre et Marie Curie</b> (Paris VI), LCP	Paris, France	GWU	<b>George Washington</b> Univ.	Washington, DC, USA
CURIN	<b>Univ. Pierre et Marie Curie</b> (Paris VI), LPNHE	Paris, France	HAHN	<b>Hahn-Meitner</b> Inst. Berlin GmbH	Berlin, Germany
CURIT	<b>Univ. Pierre et Marie Curie</b> (Paris VI), LPTHE	Paris, France	HAIF	<b>Technion</b> – Israel Inst. of Tech.	Technion, Haifa, Israel
DALH	<b>Dalhousie</b> Univ.	Halifax, NS, Canada	HAMB	Univ. <b>Hamburg</b>	Hamburg, Germany
DALI	<b>Dalian</b> Univ. of Tech.	Dalian, China	HANN	Univ. <b>Hannover</b>	Hannover, Germany
DARE	<b>Daresbury</b> Lab	Cheshire, United Kingdom	HARC	<b>Houston Advanced Research Ctr.</b>	The Woodlands, TX, USA
DARM	Tech. Hochschule <b>Darmstadt</b>	Darmstadt, Germany	HARV	<b>Harvard</b> Univ.	Cambridge, MA, USA
DELA	Univ. of <b>Delaware</b> ; Dept. of Physics & Astronomy	Newark, DE, USA	HARV	<b>Harvard</b> Univ. (LPPC)	Cambridge, MA, USA
DELH	Univ. of <b>Delhi</b>	Delhi, India	HAWA	Univ. of <b>Hawai'i</b>	Honolulu, HI, USA
DESY	<b>DESY</b> , Deutsches Elektronen-Synchrotron	Hamburg, Germany	HEBR	<b>Hebrew</b> Univ.	Jerusalem, Israel
DFAB	Escuela de Ingenieros	Bilbao, Spain	HEID	Univ. <b>Heidelberg</b> ; (unspecified division) (Historical in <i>Review</i> )	Heidelberg, Germany
DOE	<b>Department of Energy</b>	Washington, DC, USA	HEIDH	Ruprecht-Karls Univ. <b>Heidelberg</b>	Heidelberg, Germany
DORT	Technische Univ. <b>Dortmund</b>	Dortmund, Germany	HEIDP	Univ. <b>Heidelberg</b> ; Physics Inst.	Heidelberg, Germany
DUKE	<b>Duke</b> Univ.	Durham, NC, USA	HEIDT	Ruprecht-Karls-Univ. <b>Heidelberg</b>	Heidelberg, Germany
DURH	Univ. of <b>Durham</b>	Durham, United Kingdom	HELS	Univ. of <b>Helsinki</b>	University of Helsinki, Finland
DUUC	<b>University College</b> Dublin	Dublin, Ireland	HIRO	<b>Hiroshima</b> Univ.	Higashi-Hiroshima, Japan
EDIN	Univ. of <b>Edinburgh</b>	Edinburgh, United Kingdom	HOUS	Univ. of <b>Houston</b>	Houston, TX, USA
EFI	Univ. of Chicago, <b>The Enrico Fermi Inst.</b>	Chicago, IL, USA	HPC	<b>Hewlett-Packard</b> Corp.	Cupertino, CA, USA
ELMT	<b>Elmhurst</b> College	Elmhurst, IL, USA	HSCA	<b>Harvard-Smithsonian</b> Center for Astrophysics	Cambridge, MA, USA
ENSP	<b>l'Ecole Normale Supérieure</b>	Paris, France	IAS	<b>Inst. for Advanced Study</b>	Princeton, NJ, USA
EOTV	<b>Eötvös</b> University	Budapest, Hungary	IASD	<b>Dublin</b> Inst. for Advanced Studies	Dublin, Ireland
EPOL	<b>École Polytechnique</b>	Palaiseau, France	IBAR	<b>Ibaraki</b> Univ.	Ibaraki, Japan
ERLA	Univ. <b>Erlangen-Nurnberg</b>	Erlangen, Germany	IBM	<b>IBM</b> Corp.	Palo Alto, CA, USA
ETH	<b>Univ. Zürich</b>	Zürich, Switzerland	IBMY	<b>IBM</b>	Yorktown Heights, NY, USA
			IBS	Inst. for <b>Boson</b> Studies	Pasadena, CA, USA
			ICEPP	The Univ. of <b>Tokyo</b>	Tokyo, Japan
			ICRR	Univ. of <b>Tokyo</b>	Chiba, Japan

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ICTP	<b>Abdus Salam</b> International Centre for Theoretical Physics	<b>Trieste</b> , Italy	KNTY	Univ. of <b>Kentucky</b>	Lexington, KY, USA
IFIC	<b>IFIC</b> (Instituto de Física Corpuscular)	<b>Paterna (Valencia)</b> , Spain	KOBE	<b>Kobe</b> Univ.	Kobe, Japan
IFRJ	Univ. Federal do <b>Rio de Janeiro</b>	Rio de Janeiro, RJ, Brasil	KOMAB	Univ. of <b>Tokyo, Komaba</b>	Tokyo, Japan
IIT	<b>Illinois Inst. of Tech.</b>	Chicago, IL, USA	KONAN	<b>Konan</b> Univ.	Kobe, Japan
ILL	Univ. of <b>Illinois at Urbana-Champaign</b>	Urbana, IL, USA	KOSI	Inst. of Experimental Physics SAS	<b>Košice</b> , Slovakia
ILLC	Univ. of <b>Illinois at Chicago</b>	Chicago, IL, USA	KYOT	<b>Kyoto</b> Univ.; Dept. of Physics, Graduate School of Science	Kyoto, Japan
ILLG	Inst. <b>Laue-Langevin</b>	Grenoble, France	KYOTU	<b>Kyoto</b> Univ.; <b>Yukawa</b> Inst. for Theor. Physics	Kyoto, Japan
IND	<b>Indiana</b> Univ.	Bloomington, IN, USA	KYUN	<b>Kyungpook</b> National Univ.	Daegu, Republic of Korea
INEL	E G and G <b>Idaho</b> , Inc.	Idaho Falls, ID, USA	KYUSH	<b>Kyushu</b> Univ.; Elementary Particle Theory Group; Exp. Particle Physics Group; Research Center for Advanced Particle Physics	Fukuoka, Japan
INFN	Ist. Nazionale di Fisica Nucleare (Generic INFN, unknown location)	Various places, Italy	LALO	<b>LAL</b> , Laboratoire de l'Accélérateur Linéaire	<b>Orsay</b> , France
INNS	Univ. of <b>Innsbruck</b>	<b>Innsbruck</b> , Austria	LANC	<b>Lancaster</b> Univ.	Lancaster, United Kingdom
INPK	Henryk Niewodniczański Inst. of Nuclear Physics	<b>Kraków</b> , Poland	LANL	<b>Los Alamos National Lab.</b> (LANL)	Los Alamos, NM, USA
INRM	<b>INR</b> , Inst. for Nucl. Research	<b>Moscow</b> , Russian Federation	LAPL	Univ. Nacional de <b>La Plata</b>	La Plata, Argentina
INUS	<b>KEK</b> , High Energy Accelerator Research Organization	Tokyo, Japan	LAPP	<b>LAPP</b> , Lab. d'Annecy-le-Vieux de Phys. des Particules	<b>Annecy-le-Vieux</b> , France
IOAN	Univ. of <b>Ioannina</b>	Ioannina, Greece	LASL	U.C. <b>Los Alamos Scientific</b> Lab. (Old name for LANL)	Los Alamos, NM, USA
IOFF	A.F. <b>Ioffe</b> Phys. Tech. Inst.	St. <b>Petersburg</b> , Russian Federation	LATV	<b>Latvian</b> State Univ.	Riga, Latvia
IOWA	Univ. of <b>Iowa</b>	Iowa City, IA, USA	LAUS	EPFL <b>Lausanne</b>	Lausanne, Switzerland
IPN	<b>IPN</b> , Inst. de Phys. Nucl.	<b>Orsay</b> , France	LAVL	Univ. <b>Laval</b>	Quebec, QC, Canada
IPNP	Univ. Pierre et Marie Curie ( <b>Paris VI</b> )	Paris, France	LBL	<b>Lawrence Berkeley National Lab.</b>	Berkeley, CA, USA
IRAD	Inst. du Radium (Historical)	<b>Paris</b> , France	LCGT	Univ. di <b>Torino</b>	Turin, Italy
ISNG	Lab. de Physique Subatomique et de Cosmologie ( <b>LPSC</b> )	<b>Grenoble</b> , France	LEBD	<b>Lebedev</b> Physical Inst.	<b>Moscow</b> , Russian Federation
ISU	<b>Iowa State</b> Univ.	Ames, IA, USA	LECE	Univ. di <b>Lecce</b>	Lecce, Italy
ISUT	Isfahan University of Technology	Isfahan, Iran	LEED	Univ. of <b>Leeds</b>	Leeds, United Kingdom
ITEP	<b>ITEP</b> , Inst. of Theor. and Exp. Physics	<b>Moscow</b> , Russian Federation	LEGN	Lab. Naz. di Legnaro	<b>Legnaro</b> , Italy
ITHA	<b>Ithaca</b> College	Ithaca, NY, USA	LEHI	<b>Lehigh</b> Univ.	Bethlehem, PA, USA
IUPU	<b>Indiana</b> Univ., <b>Purdue</b> Univ. <b>Indianapolis</b>	Indianapolis, IN, USA	LEHM	<b>Lehman</b> College of CUNY	Bronx, NY, USA
JADA	<b>Jadavpur</b> Univ.	Calcutta, India	LEID	Univ. <b>Leiden</b>	Leiden, The Netherlands
JAGL	<b>Jagiellonian</b> Univ.	Kraków, Poland	LEMO	<b>Le Moyne</b> Coll.	Syracuse, NY, USA
JHU	<b>Johns Hopkins</b> Univ.	Baltimore, MD, USA	LEUV	Katholieke Univ. <b>Leuven</b>	Leuven, Belgium
JINR	<b>JINR</b> , Joint Inst. for Nucl. Research	<b>Dubna</b> , Russian Federation	LIEG	Univ. de <b>Liège</b>	Liège, Belgium
JULI	Forschungszentrum <b>Jülich</b>	Jülich, Germany	LINZ	Univ. <b>Linz</b>	Linz, Austria
JYV	Univ. of <b>Jyväskylä</b>	Jyväskylä, Finland	LISB	Inst. Nacional de Investigacion Cientifica	<b>Lisboa</b> CODEX, Portugal
KAGO	Univ. of <b>Kagoshima</b>	Kagoshima-shi, Japan	LISBT	Centro de Física Teórica de Partículas ( <b>CFTP</b> )	<b>Lisboa</b> , Portugal
KAIST	<b>Korea</b> Advanced Inst. of Science and Technology	Yusung ku, Daejeon, Republic of Korea	LIVP	Univ. of <b>Liverpool</b>	Liverpool, United Kingdom
KANS	Univ. of <b>Kansas</b>	Lawrence, KS, USA	LLL	<b>Lawrence Livermore Lab.</b> (Old name for LLNL)	Livermore, CA, USA
KARL	Univ. <b>Karlsruhe</b> (Historical in <i>Review</i> )	Karlsruhe, Germany	LLNL	<b>Lawrence Livermore National Lab.</b>	Livermore, CA, USA
KARLE	<b>Karlsruhe</b> Inst. of Technology (KIT); Inst. for Experimental Nuclear Physics	Karlsruhe, Germany	LOCK	<b>Lockheed</b> Palo Alto Res. Lab	Palo Alto, CA, USA
KARLK	<b>Karlsruhe</b> Inst. of Technology (KIT)	Eggenstein-Leopoldshafen, Germany	LOIC	<b>Imperial</b> College of Science Tech. & Medicine	London, United Kingdom
KARLT	<b>Karlsruhe</b> Inst. of Technology (KIT); Inst. for Theoretical Physics	Karlsruhe, Germany	LOQM	Queen Mary, <b>Univ. of London</b>	London, United Kingdom
KAZA	<b>Kazakh</b> Inst. of High Energy Physics	Alma Ata, Kazakhstan	LOUC	<b>University College</b> London	London, United Kingdom
KEK	<b>KEK</b> , High Energy Accelerator Research Organization	Ibaraki-ken, Japan	LOUV	Univ. Catholique de <b>Louvain</b>	Louvain-la-Neuve, Belgium
KENT	Univ. of <b>Kent</b>	Canterbury, United Kingdom	LOWC	<b>Westfield</b> College (Historical, see LOQM (Queen Mary and Westfield joined))	London, United Kingdom
KEYN	<b>Open</b> Univ.	Milton Keynes, United Kingdom	LRL	U.C. Lawrence Radiation Lab. (Old name for LBL)	<b>Berkeley</b> , CA, USA
KFTI	<b>Kharkov</b> Inst. of Physics and Tech. (NSC KIPT)	Kharkov, Ukraine	LSU	<b>Louisiana</b> State Univ.	Baton Rouge, LA, USA
KIAE	<b>Kurchatov</b> Inst.	<b>Moscow</b> , Russian Federation	LUND	<b>Fysiska Institutionen</b>	<b>Lund</b> , Sweden
KIAM	<b>Keldysh</b> Inst. of Applied Math., Acad. Sci., Russia	<b>Moscow</b> , Russian Federation	LUND	<b>Lund</b> Univ.	Lund, Sweden
KIDR	<b>Vinča Inst. of Nuclear Sciences</b>	Belgrade, Serbia	LYON	Institute de Physique Nucléaire de <b>Lyon</b> (IPN)	Villeurbanne, France
KIEV	<b>Institute for Nuclear Research</b>	<b>Kyiv</b> , Ukraine	MADE	<b>UAM/CSIC</b> , Inst. de Física Teórica	<b>Madrid</b> , Cantoblanco, Spain
KINK	<b>Kinki</b> Univ.	Osaka, Japan	MADR	<b>C.I.E.M.A.T</b>	<b>Madrid</b> , Spain
			MADRA	Univ. of <b>Madras</b>	Madras, India
			MADU	Univ. Autónoma de <b>Madrid</b>	Cantoblanco, Madrid, Spain
			MANI	Univ. of <b>Manitoba</b>	Winnipeg, MB, Canada

# Abbreviations Used in the Particle Listings

MANZ	<b>Johannes-Gutenberg- Univ.</b> ; Inst. für Kernphysik, J.-J.-Becher-Weg 45; Inst. für Physik, Staudingerweg 7	<b>Mainz</b> , Germany	NEAS	<b>Northeastern Univ.</b>	Boston, MA, USA
MARB	Univ. <b>Marburg</b>	Marburg, Germany	NEBR	Univ. of <b>Nebraska</b>	Lincoln, NE, USA
MARS	Centre de Physique des Par- ticules de <b>Marseille</b>	Marseille, France	NEUC	Univ. de <b>Neuchâtel</b>	Neuchâtel, Switzerland
MASA	Univ. of <b>Massachusetts</b> Amherst	<b>Amherst</b> , MA, USA	NICEA	Univ. de <b>Nice</b>	Nice, France
MASB	Univ. of <b>Massachusetts</b> Boston	<b>Boston</b> , MA, USA	NICEO	Observatoire de <b>Nice</b>	Nice, France
MASD	Univ. of <b>Massachusetts</b> Dartmouth	North <b>Dartmouth</b> , MA, USA	NIHO	<b>Nihon</b> Univ.	Tokyo, Japan
MCGI	<b>McGill</b> Univ.	Montreal, QC, Canada	NIIG	<b>Niigata</b> Univ.	Niigata, Japan
MCHS	Univ. of <b>Manchester</b>	Manchester, United Kingdom	NIJM	<b>Radboud</b> Univ. Nijmegen	AJ <b>Nijmegen</b> , The Nether- lands
MCMS	<b>McMaster</b> Univ.	Hamilton, ON, Canada	NIRS	Nat. Inst. Radiological Sci- ences	<b>Chiba</b> , Japan
MEHTA	<b>Harish-Chandra</b> Research Inst.	Allahabad, India	NIST	<b>National Institute of Stan- dards &amp; Technology</b>	Gaithersburg, MD, USA
MEIS	<b>Meisei</b> Univ.	Tokyo, Japan	NIU	<b>Northern Illinois</b> Univ.	De Kalb, IL, USA
MELB	Univ. of <b>Melbourne</b>	Victoria, Australia	NMSU	<b>New Mexico</b> State Univ.; Dept. of Physics, MSC 3D; Part. & Nucl. Phys. Group, Box 30001/Dept.	Las Cruces, NM, USA
MEUD	Observatoire de <b>Meudon</b>	Meudon, France	NORD	<b>Nordita</b>	Stockholm, Sweden
MICH	Univ. of <b>Michigan</b>	Ann Arbor, MI, USA	NOTT	Univ. of <b>Nottingham</b>	Nottingham, United Kingdom
MILA	Univ. di <b>Milano</b>	Milano, Italy	NOVM	Inst. of Mathematics	<b>Novosibirsk</b> , Russian Federa- tion
MILAI	<b>INFN</b> , Sez. di <b>Milano</b>	Milano, Italy	NOVO	BINP, <b>Budker</b> Inst. of Nu- clear Physics	<b>Novosibirsk</b> , Russian Federa- tion
MINN	Univ. of <b>Minnesota</b>	Minneapolis, MN, USA	NPOL	Polytechnic of <b>North Lon- don</b>	London, United Kingdom
MIPT	Moscow Institute of Physics and Technology	Moscow, Russian Federation	NRL	<b>Naval Research</b> Lab	Washington, DC, USA
MISS	Univ. of <b>Mississippi</b>	University, MS, USA	NSF	<b>National Science Founda- tion</b>	Arlington, VA, USA
MISSR	Univ. of <b>Missouri</b>	Rolla, MO, USA	NTHU	<b>National Tsing Hua</b> Univ.	Hsinchu, Taiwan
MIT	<b>MIT</b> <b>Massachusetts Inst.</b> <b>of Technology</b>	Cambridge, MA, USA	NTUA	<b>National Tech. Univ. of Athens</b>	Athens, Greece
MIU	<b>Maharishi</b> International Univ.	Fairfield, IA, USA	NWES	<b>Northwestern</b> Univ.	Evanston, IL, USA
MIYA	<b>Miyazaki</b> Univ.	Miyazaki-shi, Japan	NYU	<b>New York</b> Univ.	New York, NY, USA
MONP	Univ. de <b>Montpellier</b> II	Montpellier, France	OBER	<b>Oberlin</b> College	Oberlin, OH, USA
MONS	Univ. of <b>Mons</b>	<b>Mons</b> , Belgium	OCH	<b>Ochanomizu</b> Univ.	Tokyo, Japan
MONT	Univ. de <b>Montréal</b> ; Pavillon René-J.-A.-Lévesque	Montréal, PQ, Canada	OHIO	<b>Ohio</b> Univ.	Athens, OH, USA
MONTC	Univ. de <b>Montréal</b> ; Centre de recherches mathématiques	Montréal, PQ, Canada	OKAY	<b>Okayama</b> Univ.	Okayama, Japan
MOSU	<b>Skobeltsyn</b> Inst. of Nuclear Physics, Lomonosov Moscow State Univ.; Experimental HEP Division; Theoretical HEP Division	<b>Moscow</b> , Russian Federation	OKLA	Univ. of <b>Oklahoma</b>	Norman, OK, USA
MPCM	Max Planck Inst. für Chemie	<b>Mainz</b> , Germany	OKSU	<b>Oklahoma State</b> Univ.	Stillwater, OK, USA
MPEI	<b>Moscow Physical Engi- neering</b> Inst.	Moscow, Russian Federation	OREG	Univ. of <b>Oregon</b> ; Inst. of Theoretical Science; U.O. Center for High Energy Physics	Eugene, OR, USA
MPIG	<b>Max-Planck-Institute</b> für Astrophysik	Garching, Germany	ORNL	<b>Oak Ridge</b> National Labora- tory	Oak Ridge, TN, USA
MPIH	<b>Max-Planck-Inst.</b> für Kern- physik	<b>Heidelberg</b> , Germany	ORSAY	Univ. de <b>Paris Sud</b> 11	<b>Orsay</b> CEDEX, France
MPIM	<b>Max-Planck-Inst.</b> für Physik	<b>München</b> , Germany	ORST	<b>Oregon State</b> Univ.	Corvallis, OR, USA
MSST	Mississippi State University	Mississippi State, MS, USA	OSAK	<b>Osaka</b> Univ.	Osaka, Japan
MSU	<b>Michigan State</b> Univ.	East Lansing, MI, USA	OSKC	<b>Osaka City</b> Univ.	Osaka, Japan
MTHO	<b>Mount Holyoke</b> College	South Hadley, MA, USA	OSLO	Univ. of <b>Oslo</b>	Oslo, Norway
MULH	Centre Univ. du <b>Haut-Rhin</b>	Mulhouse, France	OSU	<b>Ohio State</b> Univ.	Columbus, OH, USA
MUNI	<b>Ludwig-Maximilians-Univ.</b> <b>München</b>	Garching, Germany	OTTA	Univ. of <b>Ottawa</b>	Ottawa, ON, Canada
MUNT	Tech. Univ. <b>München</b>	Garching, Germany	OXF	University of <b>Oxford</b>	Oxford, United Kingdom
MURA	<b>Midwestern</b> Univ. Research Assoc. (Historical in <i>Review</i> )	Stroughton, WI, USA	OXFTP	Univ. of <b>Oxford</b>	Oxford, United Kingdom
MURC	Univ. of Murcia	Murcia, Spain	PADO	Univ. degli Studi di <b>Padova</b>	Padova, Italy
NAAS	North Americal Aviation Sci- ence Center (Historical in <i>Review</i> )	Thousand Oaks, CA, USA	PARIN	<b>LPNHE</b> , IN <sup>2</sup> P <sup>3</sup> /CNRS	Paris, France
NAGO	<b>Nagoya</b> Univ.	Nagoya, Japan	PARIS	Univ. de Paris (Historical)	<b>Paris</b> , France
NANJ	<b>Nanjing</b> Univ.	Nanjing, China	PARIT	<b>Univ. Paris VII</b> , LP <sup>THE</sup>	Paris, France
NAPL	Univ. di <b>Napoli</b> “Federico II”	Napoli, Italy	PARM	<b>INFN</b> , Gruppo Collegato di <b>Parma</b>	Parma, Italy
NASA	<b>NASA</b>	Greenbelt, MD, USA	PAST	Institut <b>Pasteur</b>	<b>Paris</b> , France
NBS	<b>U.S National Bureau of Standards</b> (Old name for NIST)	Gaithersburg, MD, USA	PATR	Univ. of <b>Patras</b>	Patras, Greece
NBSB	<b>National Inst. Standards Tech.</b>	Boulder, CO, USA	PAVI	Univ. di <b>Pavia</b>	Pavia, Italy
NCAR	<b>National Center for Atmo- spheric Research</b>	Boulder, CO, USA	PAVII	<b>INFN</b> , Sez. di <b>Pavia</b>	<b>Pavia</b> , Italy
NCSU	<b>North Carolina</b> State Univ.	<b>Raleigh</b> , NC, USA	PENN	Univ. of <b>Pennsylvania</b>	Philadelphia, PA, USA
NDAM	Univ. of <b>Notre Dame</b>	Notre Dame, IN, USA	PGIA	<b>INFN</b> , Sezione di <b>Perugia</b>	Perugia, Italy
			PISA	Univ. di <b>Pisa</b>	Pisa, Italy
			PISAI	<b>INFN</b> , Sez. di <b>Pisa</b>	Pisa, Italy
			PITT	Univ. of <b>Pittsburgh</b>	Pittsburgh, PA, USA
			PLAT	<b>SUNY at Plattsburgh</b>	Plattsburgh, NY, USA
			PLRM	Univ. di <b>Palermo</b>	Palermo, Italy
			PNL	<b>Battelle</b> Memorial Inst.	Richland, WA, USA
			PNPI	<b>Petersburg</b> Nuclear Physics Inst. of Russian Academy of Sciences	Gatchina, Russian Federation
			PPA	Princeton-Penn. Proton Accel- erator (Historical in <i>Review</i> )	Princeton, NJ, USA



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PRAG	Inst. of Physics, <b>ASCR</b>	<b>Prague</b> , Czech Republic	SFSU	<b>California State Univ.</b>	<b>San Francisco</b> , CA, USA
PRIN	<b>Princeton Univ.</b>	Princeton, NJ, USA	SHAMS	<b>Ain Shams University</b>	Abbassia, Cairo, Egypt
PSI	Paul Scherrer Inst.	<b>Villigen</b> PSI, Switzerland	SHDN	<b>Shandong Univ.</b>	Jinan, Shandong, China
PSLL	<b>Physical Science Lab</b>	Las Cruces, NM, USA	SHEF	Univ. of <b>Sheffield</b>	Sheffield, United Kingdom
PSU	<b>Penn State Univ.</b>	University Park, PA, USA	SHMP	Univ. of <b>Southampton</b>	Southampton, United Kingdom
PUCB	<b>Pontificia Univ. Católica do Rio de Janeiro</b>	Rio de Janeiro, RJ, Brasil	SHRZ	<b>Shiraz Univ.</b>	Shiraz, Iran
PUEB	Univ. Autonoma de <b>Puebla</b>	<b>Puebla</b> , Pue, Mexico	SIEG	Univ. <b>Siegen</b>	Siegen, Germany
PURD	<b>Purdue Univ.</b>	West Lafayette, IN, USA	SILES	Univ. of <b>Silesia</b>	Katowice, Poland
QUKI	<b>Queen's Univ.</b>	Kingston, ON, Canada	SIN	Swiss Inst. of Nuclear Research (Old name for VILL)	<b>Villigen</b> , Switzerland
RAL	<b>STFC Rutherford Appleton Lab.</b>	Chilton, Didcot, Oxfordshire, United Kingdom	SING	<b>National Univ.</b> of Singapore	Kent Ridge, Singapore
REGE	Univ. <b>Regensburg</b>	Regensburg, Germany	SISSA	Scuola Internazionale Superiore di Studi Avanzati	<b>Trieste</b> , Italy
REHO	<b>Weizmann Inst.</b> of Science	Rehovot, Israel	SLAC	<b>SLAC National Accelerator Laboratory</b>	Menlo Park, CA, USA
REZ	Nuclear Physics Inst. <b>AVČR</b>	<b>Řež</b> , Czech Republic	SLOV	Inst. of Physics, Slovak Acad. of Sciences	<b>Bratislava</b> 45, Slovakia
RGSUL	Univ. Federal do <b>Rio Grande do Sul</b> (UFRGS)	Porto Alegre, RS, Brasil	SMU	<b>Southern Methodist Univ.</b>	Dallas, TX, USA
RHBL	<b>Royal Holloway, Univ. of London</b>	Egham, Surrey, United Kingdom	SNSP	<b>Scuola Normale Superiore</b>	<b>Pisa</b> , Italy
RHEL	<b>Rutherford High Energy Lab</b> (Old name for RAL)	Chilton, Didcot, Oxon., United Kingdom	SOFI	Inst. for Nuclear Research and Nuclear Energy	<b>Sofia</b> , Bulgaria
RICE	<b>Rice Univ.</b>	Houston, TX, USA	SOFU	Univ. of <b>Sofia</b> "St. Kliment Ohridski"	Sofia, Bulgaria
RIKEN	<b>Riken</b> Nishina Center for Accelerator-Based Science	Saitama, Japan	SPAUL	Univ. de <b>São Paulo</b>	São Paulo, SP, Brasil
RIKK	<b>Rikkyo Univ.</b>	Tokyo, Japan	SPIFT	Inst. de Física Teórica ( <b>IFT</b> )	<b>São Paulo</b> , SP, Brasil
RIS	<b>Rowland Inst.</b> for Science	Cambridge, MA, USA	SSL	Univ. of <b>California (Berkeley)</b>	Berkeley, CA, USA
RISC	<b>Rockwell International</b>	Thousand Oaks, CA, USA	STAN	<b>Stanford Univ.</b>	Stanford, CA, USA
RISL	<b>Universities Research Reactor</b>	<b>Risley</b> , Warrington, United Kingdom	STEV	<b>Stevens Inst.</b> of Tech.	Hoboken, NJ, USA
RISO	<b>Riso National Laboratory</b>	Roskilde, Denmark	STFN	<b>Jožef Stefan Institute</b>	<b>Ljubljana</b> , Slovenia
RL	<b>Rutherford High Energy Lab</b> (Old name for RAL)	Chilton, Didcot, Oxon., United Kingdom	STLO	<b>St. Louis Univ.</b>	St. Louis, MO, USA
RMCS	<b>Royal Military Coll.</b> of Science	Swindon, Wilts., United Kingdom	STOH	<b>Stockholm Univ.</b>	Stockholm, Sweden
ROCH	Univ. of <b>Rochester</b>	Rochester, NY, USA	STON	<b>SUNY at Stony Brook</b>	Stony Brook, NY, USA
ROCK	<b>Rockefeller Univ.</b>	New York, NY, USA	STRB	Inst. Pluridisciplinaire Hubert Curien ( <b>CNRS</b> )	<b>Strasbourg</b> , France
ROMA	Univ. di <b>Roma</b> (Historical)	<b>Roma</b> , Italy	STUT	Univ. <b>Stuttgart</b>	Stuttgart, Germany
ROMA2	Univ. di <b>Roma</b> , "Tor Vergata"	Roma, Italy	STUTM	Max-Planck-Inst.	<b>Stuttgart</b> , Germany
ROMA3	<b>INFN, Sez. di Roma Tre</b>	<b>Roma</b> , Italy	SUGI	<b>Sugiyama Jogakuen Univ.</b>	Aichi, Japan
ROMAI	<b>INFN, Sez. di Roma</b>	Roma, Italy	SURR	Univ. of <b>Surrey</b>	Guildford, Surrey, United Kingdom
ROSE	<b>Rose-Hulman Inst.</b> of Technology	Terre Haute, IN, USA	SUSS	Univ. of <b>Sussex</b>	Brighton, United Kingdom
RPI	<b>Rensselaer Polytechnic Inst.</b>	Troy, NY, USA	SVR	<b>Savannah River Labs.</b>	Aiken, SC, USA
RUTG	<b>Rutgers</b> , the State Univ. of New Jersey	Piscataway, NJ, USA	SYDN	Univ. of <b>Sydney</b>	Sydney, NSW, Australia
S0GA	<b>Sogang University</b>	Seoul, Republic of Korea	SYRA	<b>Syracuse Univ.</b>	Syracuse, NY, USA
SACL	<b>CEA Saclay</b> , IRFU	Gif-sur-Yvette, France	TAJK	Acad. Sci., Tadjik SSR	<b>Dushanbe</b> , Tadjikistan
SACL5	<b>CEA Saclay</b> – IPhT	Gif-sur-Yvette, France	TAMU	<b>Texas A&amp;M Univ.</b>	College Station, TX, USA
SACLD	<b>CEA Saclay</b> (Essonne)	Gif-sur-Yvette, France	TATA	<b>Tata Inst.</b> of Fundamental Research	Bombay, India
SAGA	<b>Saga Univ.</b>	Saga-shi, Japan	TBIL	<b>Tbilisi State University</b>	Tbilisi, Republic of Georgia
SAHA	<b>Saha Inst.</b> of Nuclear Physics	Bidhan Nagar, Calcutta, India	TELA	<b>Tel-Aviv Univ.</b>	Tel Aviv, Israel
SANG	<b>Kyoto Sangyo Univ.</b>	Kyoto-shi, Japan	TELE	<b>Teledyne Brown Engineering</b>	Huntsville, AL, USA
SANI	Ist. Superiore di Sanità	<b>Roma</b> , Italy	TEMP	<b>Temple Univ.</b>	Philadelphia, PA, USA
SASK	Univ. of <b>Saskatchewan</b>	Saskatoon, SK, Canada	TENN	Univ. of <b>Tennessee</b>	Knoxville, TN, USA
SASSO	Lab. Naz. <b>Gran Sasso</b> dell'INFN	<b>Assergi</b> (AQ), Italy	TEXA	Univ. of <b>Texas at Austin</b>	Austin, TX, USA
SAVO	Univ. de <b>Savoie</b>	Chambery, France	TGAK	<b>Tokyo Gakugei Univ.</b>	Tokyo, Japan
SBER	<b>California State Univ.</b>	<b>San Bernardino</b> , CA, USA	TGU	<b>Tohoku Gakuin Univ.</b>	Miyagi, Japan
SCHAF	W.J. <b>Schafer Assoc.</b>	Livermore, DA, USA	THES	Aristotle Univ. of <b>Thessaloniki</b> (AUTH)	Thessaloniki, Greece
SCIT	<b>Science Univ.</b> of <b>Tokyo</b>	Tokyo, Japan	TINT	<b>Tokyo Inst. of Technology</b>	Tokyo, Japan
SCOT	<b>Scottish Univ.</b> Research and Reactor Ctr.	Glasgow, United Kingdom	TISA	<b>Sagamihara Inst.</b> of Space & Astronautical Sci.	Kanagawa, Japan
SCUC	Univ. of <b>South Carolina</b>	Columbia, SC, USA	TMSK	<b>Tomsk Polytechnic Univ.</b>	<b>Tomsk</b> , Russian Federation
SEAT	<b>Seattle Pacific Coll.</b>	Seattle, WA, USA	TMTC	<b>Tokyo Metropolitan Coll.</b> Tech.	Tokyo, Japan
SEIB	Austrian Research Center, <b>Seibersdorf LTD.</b>	Seibersdorf, Austria	TMU	<b>Tokyo Metropolitan Univ.</b>	Tokyo, Japan
SEOU	<b>Korea Univ.</b> ; Dept. of Physics; HEP Group	Seoul, Republic of Korea	TNTO	Univ. of <b>Toronto</b>	Toronto, ON, Canada
SEOUL	<b>Seoul National Univ.</b> ; Center for Theoretical Physics; Dept. of Physics & Astronomy, Coll. of Natural Sciences	Seoul, Republic of Korea	TOHO	<b>Toho Univ.</b>	Chiba, Japan
SERP	<b>IHEP</b> , Inst. for High Energy Physics	Protvino, Russian Federation	TOHOK	<b>Tohoku Univ.</b>	Sendai, Japan
SETO	<b>Seton Hall Univ.</b>	South Orange, NJ, USA	TOKA	<b>Tokai Univ.</b>	Shimizu, Japan
SFLA	Univ. of <b>South Florida</b>	Tampa, FL, USA	TOKAH	<b>Tokai Univ.</b>	Hiratsuka, Japan
SFRA	<b>Simon Fraser University</b>	Burnaby, BC, Canada	TOKMS	Univ. of <b>Tokyo</b> ; Meson Science Laboratory	Tokyo, Japan
			TOKU	Univ. of <b>Tokushima</b>	Tokushima-shi, Japan
			TOKY	Univ. of <b>Tokyo</b> ; High-Energy Physics Theory Group	Tokyo, Japan
			TOKYC	Univ. of <b>Tokyo</b> ; Dept. of Chemistry	Tokyo, Japan

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TORI	Univ. degli Studi di <b>Torino</b>	Torino, Italy	UTRE	Univ. of <b>Utrecht</b>	Utrecht, The Netherlands
TPTI	<b>Uzbek</b> Academy of Sciences	<b>Tashkent</b> , Republic of Uzbekistan	UTRO	<b>Norwegian Univ. of Science &amp; Technology</b>	Trondheim, Norway
TRIN	Trinity College <b>Dublin</b>	Dublin, Ireland	UVA	Univ. of <b>Virginia</b>	Charlottesville, VA, USA
TRIU	<b>TRIUMF</b>	Vancouver, BC, Canada	UZINR	Acad. Sci., <b>Ukrainian SSR</b>	<b>Uzhgorod</b> , Ukraine
TRST	Univ. di <b>Trieste</b>	Trieste, Italy	VALE	Univ. de <b>Valencia</b>	Burjassot, <b>Valencia</b> , Spain
TRSTI	<b>INFN</b> , Sez. di <b>Trieste</b>	Trieste, Italy	VALP	<b>Valparaiso</b> Univ.	Valparaiso, IN, USA
TRSTT	Univ. degli Studi di <b>Trieste</b>	<b>Trieste</b> , Italy	VAND	<b>Vanderbilt</b> Univ.	Nashville, TN, USA
TSUK	Univ. of <b>Tsukuba</b>	Ibaraki-ken, Japan	VASS	<b>Vassar</b> College	Poughkeepsie, NY, USA
TTAM	<b>Tamagawa</b> Univ.	Tokyo, Japan	VICT	Univ. of <b>Victoria</b>	Victoria, BC, Canada
TUAT	<b>Tokyo</b> Univ. of Agriculture Tech.	Tokyo, Japan	VIEN	Inst. für Hochenergiephysik (HEPHY)	<b>Vienna</b> , Austria
TUBIN	Univ. <b>Tübingen</b>	Tübingen, Germany	VILL	<b>ETH Zürich</b>	Zürich, Switzerland
TUFTS	<b>Tufts</b> Univ.	Medford, MA, USA	VPI	<b>Virginia Tech.</b>	Blacksburg, VA, USA
TUW	<b>Technische</b> Univ. <b>Wien</b>	Vienna, Austria	VRIJ	<b>Vrije</b> Univ.	HV <b>Amsterdam</b> , The Netherlands
TUZL	<b>Tuzla</b> Univ.	Tuzla, Argentina	WABRN	Eidgenössisches Amt für Messwesen	<b>Waber</b> , Switzerland
UBA	Univ. de <b>Buenos Aires</b>	Buenos Aires, Argentina	WARS	Univ. of <b>Warsaw</b>	Warsaw, Poland
UCB	Univ. of <b>California (Berkeley)</b>	Berkeley, CA, USA	WASCR	<b>Waseda</b> Univ.; Cosmic Ray Division	Tokyo, Japan
UCD	Univ. of <b>California (Davis)</b>	Davis, CA, USA	WASH	Univ. of <b>Washington</b> ; Elem. Particle Experiment (EPE); Particle Astrophysics (PA)	Seattle, WA, USA
UCI	Univ. of <b>California (Irvine)</b>	Irvine, CA, USA	WASU	<b>Waseda</b> Univ.; Dept. of Physics, High Energy Physics Group	Tokyo, Japan
UCLA	Univ. of <b>California (Los Angeles)</b>	Los Angeles, CA, USA	WAYN	<b>Wayne</b> State Univ.	Detroit, MI, USA
UCND	<b>Union Carbide</b> Corp.	Oak Ridge, TN, USA	WESL	<b>Wesleyan</b> Univ.	Middletown, CT, USA
UCR	Univ. of <b>California (Riverside)</b>	Riverside, CA, USA	WIEN	Univ. <b>Wien</b>	Vienna, Austria
UCSB	Univ. of <b>California (Santa Barbara)</b> ; Physics Dept., High Energy Physics Experiment	Santa Barbara, CA, USA	WILL	Coll. of <b>William and Mary</b>	Williamsburg, VA, USA
UCSBT	Univ. of <b>California (Santa Barbara)</b> ; Kavli Inst. for Theoretical Physics	Santa Barbara, CA, USA	WINR	National Centre for Nuclear Research	<b>Warsaw</b> , Poland
UCSC	Univ. of <b>California (Santa Cruz)</b>	Santa Cruz, CA, USA	WISC	Univ. of <b>Wisconsin</b>	Madison, WI, USA
UCSD	Univ. of <b>California (San Diego)</b>	La Jolla, CA, USA	WITW	Univ. of the <b>Witwatersrand</b>	Wits, South Africa
UGAZ	Univ. of <b>Gaziantep</b>	Gaziantep, Turkey	WMIU	<b>Western Michigan</b> Univ.	Kalamazoo, MI, USA
UMD	Univ. of <b>Maryland</b>	College Park, MD, USA	WONT	The Univ. of <b>Western Ontario</b>	London, ON, Canada
UNAM	Univ. Nac. Autónoma de México ( <b>UNAM</b> )	<b>México</b> , DF, Mexico	WOOD	<b>Woodstock</b> College (No longer in existence)	Woodstock, MD, USA
UNAM	Univ. Nacional Autónoma de México ( <b>UNAM</b> )	<b>México</b> , DF, Mexico	WUPP	<b>Bergische</b> Univ. Wuppertal	<b>Wuppertal</b> , Germany
UNC	Univ. of <b>North Carolina</b>	Greensboro, NC, USA	WURZ	Univ. <b>Würzburg</b>	Würzburg, Germany
UNCCH	Univ. of <b>North Carolina</b> at Chapel Hill	Chapel Hill, NC, USA	WUSL	<b>Washington</b> Univ.	St. Louis, MO, USA
UNCS	<b>Union</b> College	Schenectady, NY, USA	WYOM	Univ. of <b>Wyoming</b>	Laramie, WY, USA
UNESP	UNESP	Botucatu, Brasil	YALE	<b>Yale</b> Univ.	New Haven, CT, USA
UNH	Univ. of <b>New Hampshire</b>	Durham, NH, USA	YARO	<b>Yaroslavl</b> State Univ.	Yaroslavl, Russian Federation
UNM	Univ. of <b>New Mexico</b>	Albuquerque, NM, USA	YCC	<b>Yokohama</b> Coll. of Commerce	Yokohama, Japan
UOEH	Univ. of Occupational and Environmental Health	<b>Kitakyushu</b> , Japan	YERE	<b>Yerevan</b> Physics Inst.	Yerevan, Armenia
UPNJ	<b>Uppsala</b> College	East Orange, NJ, USA	YOKO	<b>Yokohama National</b> Univ.	Yokohama-shi, Japan
UPPS	<b>Uppsala</b> Univ.	<b>Uppsala</b> , Sweden	YORKC	<b>York</b> Univ.	Toronto, Canada
UPR	Univ. of <b>Puerto Rico</b>	<b>San Juan</b> , PR, USA	ZAGR	<b>Zagreb</b> Univ.	Zagreb, Croatia
URI	Univ. of <b>Rhode Island</b>	Kingston, RI, USA	ZARA	Univ. de <b>Zaragoza</b>	Zaragoza, Spain
USC	Univ. of <b>Southern California</b>	Los Angeles, CA, USA	ZEEM	Univ. van <b>Amsterdam</b>	TV Amsterdam, The Netherlands
USF	Univ. of <b>San Francisco</b>	San Francisco, CA, USA	ZHON	<b>Zhongshan</b> (Sun Yat-Sen) Univ.	Guangzhou, China
UTAH	Univ. of <b>Utah</b>	Salt Lake City, UT, USA	ZHZH	<b>Zhengzhou</b> Univ.	Zhengzhou, Henan, China
			ZURI	Univ. <b>Zürich</b>	Zürich, Switzerland

## GAUGE AND HIGGS BOSONS

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See key on page 601

## Gauge &amp; Higgs Boson Particle Listings

 $\gamma$ 

## GAUGE AND HIGGS BOSONS

 $\gamma$  (photon)

$$I(J^{PC}) = 0.1(1^{--})$$

 $\gamma$  MASS

Results prior to 2008 are critiqued in GOLDHABER 10. All experimental results published prior to 2005 are summarized in detail by TU 05.

The following conversions are useful:  $1 \text{ eV} = 1.783 \times 10^{-33} \text{ g} = 1.957 \times 10^{-6} m_e$ ;  $\lambda_C = (1.973 \times 10^{-7} \text{ m}) \times (1 \text{ eV}/m_\gamma)$ .

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1 <math>\times 10^{-18}</math></b>		<b>1 RYUTOV 07</b>		MHD of solar wind
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.3 $\times 10^{-9}$	95	2 EGOROV 14	COSM	Lensed quasar position
		3 ACCIOLY 10		Anomalous mag. mom.
<1 $\times 10^{-26}$		4 ADELBERGER 07A		Proca galactic field
no limit feasible		4 ADELBERGER 07A		$\gamma$ as Higgs particle
<1 $\times 10^{-19}$		5 TU 06		Torque on rotating magnetized toroid
<1.4 $\times 10^{-7}$		ACCIOLY 04		Dispersion of GHz radio waves by sun
<2 $\times 10^{-16}$		6 FULLEKRUG 04		Speed of 5-50 Hz radiation in atmosphere
<7 $\times 10^{-19}$		7 LUO 03		Torque on rotating magnetized toroid
<1 $\times 10^{-17}$		8 LAKES 98		Torque on toroid balance
<6 $\times 10^{-17}$		9 RYUTOV 97		MHD of solar wind
<8 $\times 10^{-16}$		10 FISCHBACH 94		Earth magnetic field
<5 $\times 10^{-13}$		11 CHERNIKOV 92	SQID	Ampere-law null test
<1.5 $\times 10^{-9}$	90	12 RYAN 85		Coulomb-law null test
<3 $\times 10^{-27}$		13 CHIBISOV 76		Galactic magnetic field
<6 $\times 10^{-16}$	99.7	14 DAVIS 75		Jupiter magnetic field
<7.3 $\times 10^{-16}$		HOLLWEG 74		Alfvén waves
<6 $\times 10^{-17}$		15 FRANKEN 71		Low freq. res. cir.
<2.4 $\times 10^{-13}$		16 KROLL 71A		Dispersion in atmosphere
<1 $\times 10^{-14}$		17 WILLIAMS 71	CNTR	Tests Gauss law
<2.3 $\times 10^{-15}$		GOLDHABER 68		Satellite data

- 1 RYUTOV 07 extends the method of RYUTOV 97 to the radius of Pluto's orbit.
- 2 EGOROV 14 studies chromatic dispersion of lensed quasar positions ("gravitational rainbows") that could be produced by any of several mechanisms, among them via photon mass. Limit not competitive but obtained on cosmological distance scales.
- 3 ACCIOLY 10 limits come from possible alterations of anomalous magnetic moment of electron and gravitational deflection of electromagnetic radiation. Reported limits are not "claimed" by the authors and in any case are not competitive.
- 4 When trying to measure  $m$  one must distinguish between measurements performed on large and small scales. If the photon acquires mass by the Higgs mechanism, the large-scale behavior of the photon might be effectively Maxwellian. If, on the other hand, one postulates the Proca regime for all scales, the very existence of the galactic field implies  $m < 10^{-26} \text{ eV}$ , as correctly calculated by YAMAGUCHI 59 and CHIBISOV 76.
- 5 TU 06 continues the work of LUO 03, with extended LAKES 98 method, reporting the improved limit  $\mu^2 A = (0.7 \pm 1.7) \times 10^{-13} \text{ T/m}$  if  $A = 0.2 \mu\text{G}$  out to  $4 \times 10^{22} \text{ m}$ . Reported result  $\mu = (0.9 \pm 1.5) \times 10^{-52} \text{ g}$  reduces to the frequentist mass limit  $1.2 \times 10^{-19} \text{ eV}$  (FELDMAN 98).
- 6 FULLEKRUG 04 adopted KROLL 71A method with newer and better Schumann resonance data. Result questionable because assumed frequency shift with photon mass is assumed to be linear. It is quadratic according to theorem by GOLDHABER 71B, KROLL 71, and PARK 71.
- 7 LUO 03 extends LAKES 98 technique to set a limit on  $\mu^2 A$ , where  $\mu^{-1}$  is the Compton wavelength  $\lambda_C$  of the massive photon and  $A$  is the ambient vector potential. The important departure is that the apparatus rotates, removing sensitivity to the direction of  $A$ . They take  $A = 10^{12} \text{ Tm}$ , due to "cluster level fields." But see comment of GOLDHABER 03 and reply by LUO 03b.
- 8 LAKES 98 reports limits on torque on a toroid Cavendish balance, obtaining a limit on  $\mu^2 A < 2 \times 10^{-9} \text{ Tm}^2/\text{m}^2$  via the Maxwell-Proca equations, where  $\mu^{-1}$  is the characteristic length associated with the photon mass and  $A$  is the ambient vector potential in the Lorentz gauge. Assuming  $A \approx 1 \times 10^{12} \text{ Tm}$  due to cluster fields he obtains  $\mu^{-1} > 2 \times 10^{10} \text{ m}$ , corresponding to  $\mu < 1 \times 10^{-17} \text{ eV}$ . A more conservative limit, using  $A \approx (1 \mu\text{G}) \times (600 \text{ pc})$  based on the galactic field, is  $\mu^{-1} > 1 \times 10^9 \text{ m}$  or  $\mu < 2 \times 10^{-16} \text{ eV}$ .
- 9 RYUTOV 97 uses a magnetohydrodynamics argument concerning survival of the Sun's field to the radius of the Earth's orbit. "To reconcile observations to theory, one has to reduce [the photon mass] by approximately an order of magnitude compared with" per DAVIS 75. "Secure limit, best by this method" (per GOLDHABER 10).
- 10 FISCHBACH 94 analysis is based on terrestrial magnetic fields; approach analogous to DAVIS 75. Similar result based on a much smaller planet probably follows from more precise  $B$  field mapping. "Secure limit, best by this method" (per GOLDHABER 10).
- 11 CHERNIKOV 92, motivated by possibility that photon exhibits mass only below some unknown critical temperature, searches for departure from Ampere's Law at  $1.24 \text{ K}$ . See also RYAN 85.
- 12 RYAN 85, motivated by possibility that photon exhibits mass only below some unknown critical temperature, sets mass limit at  $< (1.5 \pm 1.4) \times 10^{-42} \text{ g}$  based on Coulomb's Law departure limit at  $1.36 \text{ K}$ . We report the result as frequentist 90% CL (FELDMAN 98).
- 13 CHIBISOV 76 depends in critical way on assumptions such as applicability of virial theorem. Some of the arguments given only in unpublished references.

- 14 DAVIS 75 analysis of Pioneer-10 data on Jupiter's magnetic field. "Secure limit, best by this method" (per GOLDHABER 10).
- 15 FRANKEN 71 method is of dubious validity (KROLL 71A, JACKSON 99, GOLDHABER 10, and references therein).
- 16 KROLL 71A used low frequency Schumann resonances in cavity between the conducting earth and resistive ionosphere, overcoming objections to resonant-cavity methods (JACKSON 99, GOLDHABER 10, and references therein). "Secure limit, best by this method" (per GOLDHABER 10).
- 17 WILLIAMS 71 is landmark test of Coulomb's law. "Secure limit, best by this method" (per GOLDHABER 10).

 $\gamma$  CHARGE

OKUN 06 has argued that schemes in which all photons are charged are inconsistent. He says that if a neutral photon is also admitted to avoid this problem, then other problems emerge, such as those connected with the emission and absorption of charged photons by charged particles. He concludes that in the absence of a self-consistent phenomenological basis, interpretation of experimental data is at best difficult.

VALUE (e)	CHARGE	DOCUMENT ID	TECN	COMMENT
<b>&lt;1 <math>\times 10^{-46}</math> mixed</b>		<b>1 ALTSCHUL 07b</b>	VLBI	Aharonov-Bohm effect
<b>&lt;1 <math>\times 10^{-35}</math> single</b>		<b>2 CAPRINI 05</b>	CMB	Isotropy constraint
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1 $\times 10^{-32}$	single	1 ALTSCHUL 07b	VLBI	Aharonov-Bohm effect
<3 $\times 10^{-33}$	mixed	3 KOBYCHEV 05	VLBI	Smear as function of $B \cdot E_\gamma$
<4 $\times 10^{-31}$	single	3 KOBYCHEV 05	VLBI	Deflection as function of $B \cdot E_\gamma$
<8.5 $\times 10^{-17}$		4 SEMERTZIDIS 03		Laser light deflection in B-field
<3 $\times 10^{-28}$	single	5 SIVARAM 95	CMB	For $\Omega_M = 0.3$ , $h^2 = 0.5$
<5 $\times 10^{-30}$		6 RAFFELT 92	TOF	Pulsar $f_1 - f_2$
<2 $\times 10^{-28}$		7 COCCONI 92		VLBA radio telescope resolution
<2 $\times 10^{-32}$		COCCONI 88	TOF	Pulsar $f_1 - f_2$ TOF

- 1 ALTSCHUL 07b looks for Aharonov-Bohm phase shift in addition to geometric phase shift in radio interference fringes (VSOP mission).
- 2 CAPRINI 05 uses isotropy of the cosmic microwave background to place stringent limits on possible charge asymmetry of the Universe. Charge limits are set on the photon, neutrino, and dark matter particles. Valid if charge asymmetries produced by different particles are not anticorrelated.
- 3 KOBYCHEV 05 considers a variety of observable effects of photon charge for extragalactic compact radio sources. Best limits if source observed through a foreground cluster of galaxies.
- 4 SEMERTZIDIS 03 reports the first laboratory limit on the photon charge in the last 30 years. Straightforward improvements in the apparatus could attain a sensitivity of  $10^{-20} \text{ e}$ .
- 5 SIVARAM 95 requires that CMB photon charge density not overwhelm gravity. Result scales as  $\Omega_M h^2$ .
- 6 RAFFELT 94 notes that COCCONI 88 neglects the fact that the time delay due to dispersion by free electrons in the interstellar medium has the same photon energy dependence as that due to bending of a charged photon in the magnetic field. His limit is based on the assumption that the entire observed dispersion is due to photon charge. It is a factor of 200 less stringent than the COCCONI 88 limit.
- 7 See COCCONI 92 for less stringent limits in other frequency ranges. Also see RAFFELT 94 note.

 $\gamma$  REFERENCES

EGOROV 14	MNRAS 437 L90	P. Egorov et al.	(MOSU, MIPT, INRM)
ACCIOLY 10	PR D82 065026	A. Accioly, J. Helayel-Neto, E. Scatena	(LABEX+)
GOLDHABER 10	RMP 82 939	A.S. Goldhaber, M.M. Nieto	(STON, LANL)
ADELBERGER 07A	PRL 98 010402	E. Adelberger, G. Dvali, A. Gruzinov	(WASH, NYU)
ALTSCHUL 07b	PRL 98 261801	B. Altschul	(IND)
Also	ASP 29 280	B. Altschul	(SUCU)
RYUTOV 07	PPCF 49 B429	D.D. Ryutov	(LLNL)
OKUN 06	APP B37 565	L.B. Okun	(ITEP)
TU 06	PL A352 267	L.-C. Tu et al.	
CAPRINI 05	JCAP 0502 006	C. Caprini, P.G. Ferreira	(GEVA, OXFTP)
KOBYCHEV 05	AL 31 147	V.V. Kobychiev, S.B. Popov	(KIEV, PADO)
TU 05	RPP 68 77	L.-C. Tu, J. Luo, G.T. Gillies	
ACCIOLY 04	PR D69 107501	A. Accioly, R. Paszko	
FULLEKRUG 04	PRL 93 043901	M. Fullekrug	
GOLDHABER 03	PRL 91 149101	A.S. Goldhaber, M.M. Nieto	
LUO 03	PRL 90 081801	J. Luo et al.	
LUO 03b	PRL 91 149102	J. Luo et al.	
SEMERTZIDIS 03	PR D67 017701	N.K. Semertzidis, G.T. Danby, D.M. Lazarus	(3rd ed., J. Wiley and Sons (1999))
JACKSON 99	Classical Electrodynamics	J.D. Jackson	
FELDMAN 98	PR D57 3873	G.J. Feldman, R.D. Cousins	
LAKES 98	PRL 80 1826	R. Lakes	(WISC)
RYUTOV 97	PPCF 39 A73	D.D. Ryutov	(LLNL)
SIVARAM 95	AJP 63 473	C. Sivaram	(BANG)
FISCHBACH 94	PRL 73 514	E. Fischbach et al.	(PURD, JHU+)
RAFFELT 94	PR D50 7729	G. Raffelt	(MPIM)
CHERNIKOV 92	PRL 68 3383	M.A. Chernikov et al.	(ETH)
Also	PRL 69 2939 (erratum)	M.A. Chernikov et al.	(ETH)
COCCONI 92	AJP 60 750	G. Cocconi	(CERN)
COCCONI 88	PL B206 705	G. Cocconi	(CERN)
RYAN 85	PR D32 802	J.J. RYAN, F. Accetta, R.H. Austin	(PRIN)
CHIBISOV 76	SPU 19 624	G.V. Chibisov	(LEBD)
	Translated from UFN 119 551		
DAVIS 75	PRL 35 1402	L. Davis, A.S. Goldhaber, M.M. Nieto	(CIT, STON+)
HOLLWEG 74	PRL 32 961	J.V. Hollweg	(NCAR)
FRANKEN 71	PRL 26 115	P.A. Franken, G.W. Ampulski	(MICH)
GOLDHABER 71B	RMP 43 277	A.S. Goldhaber, M.M. Nieto	(STON, BOHR, UCSB)
KROLL 71	PRL 26 1395	N.M. Kroll	(SLAC)
KROLL 71A	PRL 27 340	N.M. Kroll	(SLAC)
PARK 71	PRL 26 1393	D. Park, E.R. Williams	(WILC)
WILLIAMS 71	PRL 26 721	E.R. Williams, J.E. Faller, H.A. Hill	(WESL)
GOLDHABER 68	PRL 21 567	A.S. Goldhaber, M.M. Nieto	(STON)
YAMAGUCHI 59	PTPS 11 37	Y. Yamaguchi	

# Gauge & Higgs Boson Particle Listings

## *g*, graviton, *W*

***g***  
or gluon

SU(3) color octet  
Mass *m* = 0. Theoretical value. A mass as large as a few MeV may not be precluded, see YNDURAIN 95.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	ABREU 92E	DLPH	Spin 1, not 0
	ALEXANDER 91H	OPAL	Spin 1, not 0
	BEHREND 82D	CELL	Spin 1, not 0
	BERGER 80D	PLUT	Spin 1, not 0
	BRANDELIK 80C	TASS	Spin 1, not 0

### gluon REFERENCES

YNDURAIN 95	PL B345 524	F.J. Yndurain	(MADU)
ABREU 92E	PL B274 498	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ALEXANDER 91H	ZPHY C52 543	G. Alexander <i>et al.</i>	(OPAL Collab.)
BEHREND 82D	PL B110 329	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BERGER 80D	PL B97 459	C. Berger <i>et al.</i>	(PLUTO Collab.)
BRANDELIK 80C	PL B97 453	R. Brandelik <i>et al.</i>	(TASSO Collab.)

**graviton**

$J = 2$

### graviton MASS

In 1970 van Dam amd Veltman (VANDAM 70) showed that "... there is a discrete difference between the theory with zero-mass and a theory with finite mass, no matter how small as compared to all external momenta. ... We may conclude that the graviton has rigorously zero mass." However, see GOLDHABER 10 and references therein. It has been of interest to set experimental limits, whether or not a finite mass can exist. In most (but not all) cases limits have been set on the distance without evidence for a Yukawa cutoff. *h*<sub>0</sub> is the Hubble constant in units of 100 km s<sup>-1</sup> Mpc<sup>-1</sup>.

The following conversions are useful: 1 eV = 1.783 × 10<sup>-33</sup> g = 1.957 × 10<sup>-6</sup> *m<sub>e</sub>*;  $\lambda_C = (1.973 \times 10^{-7} \text{ m}) \times (1 \text{ eV} / m_g)$ .

VALUE (eV)	DOCUMENT ID	COMMENT
<6 × 10 <sup>-32</sup>	1 CHOUDHURY 04	Weak gravitational lensing
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<1.2 × 10 <sup>-22</sup>	2 ABBOTT 16	LIGO black holes merger
<5 × 10 <sup>-23</sup>	3 BRITO 13	Spinning black holes bounds
<4 × 10 <sup>-25</sup>	4 BASKARAN 08	Graviton phase velocity fluctuations
<6 × 10 <sup>-32</sup>	5 GRUZINOV 05	Solar System observations
<9.0 × 10 <sup>-34</sup>	6 GERSHTEIN 04	From $\Omega_{tot}$ value assuming RTG
>6 × 10 <sup>-34</sup>	7 DVALI 03	Horizon scales
<8 × 10 <sup>-20</sup>	8,9 FINN 02	Binary pulsar orbital period decrease
	9,10 DAMOUR 91	Binary pulsar PSR 1913+16
<2 × 10 <sup>-29</sup> <i>h</i> <sub>0</sub> <sup>-1</sup>	GOLDHABER 74	Rich clusters
<7 × 10 <sup>-28</sup>	HARE 73	Galaxy
<8 × 10 <sup>4</sup>	HARE 73	2 $\gamma$ decay

- 1 CHOUDHURY 04 concludes from a study of weak-lensing data that masses heavier than about the inverse of 100 Mpc seem to be ruled out if the gravitation field has the Yukawa form.
- 2 ABBOTT 16 assumes modified dispersion relation for gravitational waves.
- 3 BRITO 13 explore massive graviton (spin-2) fluctuations around rotating black holes.
- 4 BASKARAN 08 consider fluctuations in pulsar timing due to photon interactions ("surfing") with background gravitational waves.
- 5 GRUZINOV 05 uses the DGP model (DVALI 00) showing that non-perturbative effects restore continuity with Einstein's equations as the graviton mass approaches 0, then bases his limit on Solar System observations.
- 6 GERSHTEIN 04 use non-Einstein field relativistic theory of gravity (RTG), with a massive graviton, to obtain the 95% CL mass limit implied by the value of  $\Omega_{tot} = 1.02 \pm 0.02$  current at the time of publication.
- 7 DVALI 03 suggest scale of horizon distance via DGP model (DVALI 00). For a horizon distance of  $3 \times 10^{26}$  m (about age of Universe/*c*; GOLDHABER 10) this graviton mass limit is implied.
- 8 FINN 02 analyze the orbital decay rates of PSR B1913+16 and PSR B1534+12 with a possible graviton mass as a parameter. The combined frequentist mass limit is at 90%CL.
- 9 As of 2014, limits on d*P*/d*t* are now about 0.1% (see T. Damour, "Experimental tests of gravitational theory," in this Review).
- 10 DAMOUR 91 is an analysis of the orbital period change in binary pulsar PSR 1913+16, and confirms the general relativity prediction to 0.8%. "The theoretical importance of the [rate of orbital period decay] measurement has long been recognized as a direct confirmation that the gravitational interaction propagates with velocity *c* (which is the immediate cause of the appearance of a damping force in the binary pulsar system) and thereby as a test of the existence of gravitational radiation and of its quadrupolar nature." TAYLOR 93 adds that orbital parameter studies now agree with general relativity to 0.5%, and set limits on the level of scalar contribution in the context of a family of tensor [spin 2]-biscalar theories.

### graviton REFERENCES

ABBOTT 16	PRL 116 061102	B.P. Abbott <i>et al.</i>	(LIGO and Virgo Collabs.)
BRITO 13	PR D88 023514	R. Brito, V. Cardoso, P. Pani	(LISB, MISS, HSCA+)
GOLDHABER 10	RMP 82 939	A.S. Goldhaber, M.M. Nieto	(STON, LANL)
BASKARAN 08	PR D78 044018	D. Baskaran <i>et al.</i>	
GRUZINOV 05	NAST 10 311	A. Gruzinov	(NYU)
CHOUDHURY 04	ASP 21 559	S.R. Choudhury <i>et al.</i>	(DELPH, MELB)
GERSHTEIN 04	PAN 67 1596	S.S. Gershtein <i>et al.</i>	(SERP)
	Translated from YAF 67 1618.		
DVALI 03	PR D68 024012	G.R. Dvali, A. Gruzinov, M. Zaldarriaga	(NYU)
FINN 02	PR D65 044022	L.S. Finn, P.J. Sutton	
DVALI 00	PL B485 208	G.R. Dvali, G. Gabadadze, M. Porrati	(NYU)
TAYLOR 93	NAT 355 132	J.N. Taylor <i>et al.</i>	(PRIN, ARCBO, BURE+)
DAMOUR 91	APJ 366 501	T. Damour, J.H. Taylor	(BURE, MEUD, PRIN)
GOLDHABER 74	PR D9 1119	A.S. Goldhaber, M.M. Nieto	(LANL, STON)
HARE 73	CJP 51 431	M.G. Hare	(SASK)
VANDAM 70	NP B22 397	H. van Dam, M. Veltman	(UTRE)

***W***

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## THE MASS AND WIDTH OF THE W BOSON

Revised September 2013 by M.W. Gr newald (U. College Dublin and U. Ghent) and A. Gurtu (Formerly Tata Inst.).

Precision determination of the W-mass is of great importance in testing the internal consistency of the Standard Model. From the time of its discovery in 1983, the W-boson has been studied and its mass determined in p  and e<sup>+</sup>e<sup>-</sup> interactions; it is currently studied in pp interactions at the LHC. The W mass and width definition used here corresponds to a Breit-Wigner with mass-dependent width.

Production of on-shell W bosons at hadron colliders is tagged by the high *p<sub>T</sub>* charged lepton from its decay. Owing to the unknown parton-parton effective energy and missing energy in the longitudinal direction, the collider experiments reconstruct the transverse mass of the W, and derive the W mass from comparing the transverse mass distribution with Monte Carlo predictions as a function of *M<sub>W</sub>*. These analyses use the electron and muon decay modes of the W boson.

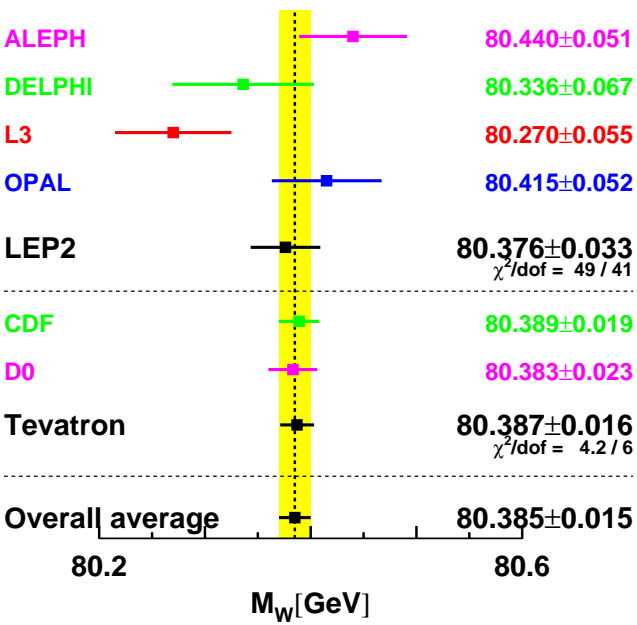
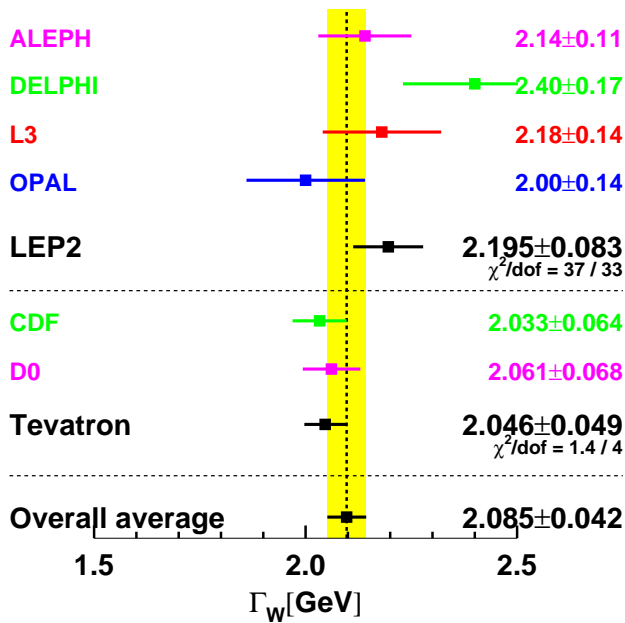


Figure 1: Measurements of the W-boson mass by the LEP and Tevatron experiments.

In the  $e^+e^-$  collider (LEP) a precise knowledge of the beam energy enables one to determine the  $e^+e^- \rightarrow W^+W^-$  cross section as a function of center of mass energy, as well as to reconstruct the W mass precisely from its decay products, even if one of them decays leptonically. Close to the  $W^+W^-$  threshold (161 GeV), the dependence of the W-pair production cross section on  $M_W$  is large, and this was used to determine  $M_W$ . At higher energies (172 to 209 GeV) this dependence is much weaker and W-bosons were directly reconstructed and the mass determined as the invariant mass of its decay products, improving the resolution with a kinematic fit.



**Figure 2:** Measurements of the W-boson width by the LEP and Tevatron experiments.

In order to compute the LEP average W mass, each experiment provided its measured W mass for the  $q\bar{q}q\bar{q}$  and  $q\bar{q}\ell\bar{\nu}_\ell$ ,  $\ell = e, \mu, \tau$  channels at each center-of-mass energy, along with a detailed break-up of errors: statistical, uncorrelated, partially correlated and fully correlated systematics [1]. These have been combined to obtain a LEP W mass of  $M_W = 80.376 \pm 0.033$  GeV. Errors due to uncertainties in LEP energy (9 MeV), and possible effect of color reconnection (CR) and Bose-Einstein correlations (BEC) between quarks from different W's (8 MeV) are included. The mass difference between  $q\bar{q}q\bar{q}$  and  $q\bar{q}\ell\bar{\nu}_\ell$  final states (due to possible CR and BEC effects) is  $-12 \pm 45$  MeV. In a similar manner, the width results obtained at LEP have been combined, resulting in  $\Gamma_W = 2.195 \pm 0.083$  GeV [1].

The two Tevatron experiments have also identified common systematic errors. Between the two experiments, uncertainties due to the parton distribution functions, radiative corrections, and choice of mass (width) in the width (mass) measurements are treated as correlated. An average W width

of  $\Gamma_W = 2.046 \pm 0.049$  GeV [2] is obtained. Errors of 20 MeV and 7 MeV accounting for PDF and radiative correction uncertainties in this width combination dominate the correlated uncertainties. At the 2012 winter conferences, the CDF and D0 experiments have presented new results for the mass of the W boson based on  $2-4 \text{ fb}^{-1}$  of Run-II data,  $80.387 \pm 0.019$  GeV [3] and  $80.375 \pm 0.023$  GeV [4], respectively. The W-mass determination from the Tevatron experiments has thus become very precise. Combining all Tevatron results from Run-I and Run-II using an improved treatment of correlations, a new average of  $80.387 \pm 0.016$  GeV is obtained [5], with common uncertainties of 10 MeV (PDF) and 4 MeV (radiative corrections).

The LEP and Tevatron results on mass and width, which are based on all results available, are compared in Fig. 1 and Fig. 2. Good agreement between the results is observed. Combining these results, assuming no common systematic uncertainties between the LEP and the Tevatron measurements, yields an average W mass of  $M_W = 80.385 \pm 0.015$  GeV and a W width of  $\Gamma_W = 2.085 \pm 0.042$  GeV.

The Standard Model prediction from the electroweak fit, using Z-pole data plus  $m_{\text{top}}$  measurement, gives a W-boson mass of  $M_W = 80.363 \pm 0.020$  GeV and a W-boson width of  $\Gamma_W = 2.091 \pm 0.002$  GeV [1].

## References

1. The LEP Collaborations: ALEPH, DELPHI, L3, OPAL, the LEP Electroweak Working Group, CERN-PH-EP/2013-022, [arXiv:1302.3415 \[hep-ex\]](#), Phys.Rept. 532 (2013) 119-244.
2. The Tevatron Electroweak Working Group, for the CDF and D0 Collaborations: *Combination of CDF and D0 Results on the Width of the W Boson*, March 2010, [arXiv:1003.2826 \[hep-ex\]](#).
3. The CDF Collaboration, *Precise measurement of the W-boson mass with the CDF II detector*, [arXiv:1203.0275 \[hep-ex\]](#), Phys. Rev. Lett. 108, 151803 (2012).
4. The D0 Collaboration, *Measurement of the W Boson Mass with the D0 Detector*, [arXiv:1203.0293 \[hep-ex\]](#), Phys. Rev. Lett. 108, 151804 (2012).
5. The CDF and D0 Collabs: *Combination of CDF and D0 W-Boson Mass Measurements*, July 2013, [arXiv:1307.7627 \[hep-ex\]](#), Phys. Rev. D88, 052018 (2013).

## W MASS

The W-mass listed here corresponds to the mass parameter in a Breit-Wigner distribution with mass-dependent width. To obtain the world average, common systematic uncertainties between experiments are properly taken into account. The LEP-2 average W mass based on published results is  $80.376 \pm 0.033$  GeV [SCHAE13a]. The combined Tevatron data yields an average W mass of  $80.387 \pm 0.016$  GeV [AALTONEN13a].

OUR FIT uses these average LEP and Tevatron mass values and combines them assuming no correlations.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>80.385 ± 0.015 OUR FIT</b>				
80.375 ± 0.023	2177k	1 ABAZOV	14N D0	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
80.387 ± 0.019	1095k	2 AALTONEN	12E CDF	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
80.336 ± 0.055 ± 0.039	10.3k	3 ABDALLAH	08A DLPH	$E_{\text{cm}}^{e^+e^-} = 161-209 \text{ GeV}$
80.415 ± 0.042 ± 0.031	11830	4 ABBIENDI	06 OPAL	$E_{\text{cm}}^{e^+e^-} = 170-209 \text{ GeV}$
80.270 ± 0.046 ± 0.031	9909	5 ACHARD	06 L3	$E_{\text{cm}}^{e^+e^-} = 161-209 \text{ GeV}$
80.440 ± 0.043 ± 0.027	8692	6 SCHAE1	06 ALEP	$E_{\text{cm}}^{e^+e^-} = 161-209 \text{ GeV}$
80.483 ± 0.084	49247	7 ABAZOV	02D D0	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
80.433 ± 0.079	53841	8 AFFOLDER	01E CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$

# Gauge & Higgs Boson Particle Listings

## W

• • • We do not use the following data for averages, fits, limits, etc. • • •

80.367 ± 0.026	1677k	<sup>9</sup> ABAZOV	12F	D0	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
80.401 ± 0.043	500k	<sup>10</sup> ABAZOV	09AB	D0	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
80.413 ± 0.034 ± 0.034	115k	<sup>11</sup> AALTONEN	07F	CDF	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
82.87 ± 1.82 $^{+0.30}_{-0.16}$	1500	<sup>12</sup> AKTAS	06	H1	$e^{\pm} p \rightarrow \bar{\nu}_e(\nu_e) X, \sqrt{s} \approx 300 \text{ GeV}$
80.3 ± 2.1 ± 1.2 ± 1.0	645	<sup>13</sup> CHEKANOV	02c	ZEUS	$e^- p \rightarrow \nu_e X, \sqrt{s} = 318 \text{ GeV}$
81.4 $^{+2.7}_{-2.6}$ ± 2.0 $^{+3.3}_{-3.0}$	1086	<sup>14</sup> BREITWEG	00D	ZEUS	$e^+ p \rightarrow \bar{\nu}_e X, \sqrt{s} \approx 300 \text{ GeV}$
80.84 ± 0.22 ± 0.83	2065	<sup>15</sup> ALITTI	92B	UA2	See $W/Z$ ratio below
80.79 ± 0.31 ± 0.84		<sup>16</sup> ALITTI	90B	UA2	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
80.0 ± 3.3 ± 2.4	22	<sup>17</sup> ABE	89i	CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
82.7 ± 1.0 ± 2.7	149	<sup>18</sup> ALBAJAR	89	UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
81.8 $^{+6.0}_{-5.3}$ ± 2.6	46	<sup>19</sup> ALBAJAR	89	UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
89 ± 3 ± 6	32	<sup>20</sup> ALBAJAR	89	UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
81. ± 5.	6	ARNISON	83	UA1	$E_{\text{cm}}^{ee} = 546 \text{ GeV}$
80. $^{+10.}_{-6.}$	4	BANNER	83B	UA2	Repl. by ALITTI 90B

<sup>1</sup> ABAZOV 14N is a combination of ABAZOV 09AB and ABAZOV 12F, also giving more details on the analysis.

<sup>2</sup> AALTONEN 12E select 470k  $W \rightarrow e\nu$  decays and 625k  $W \rightarrow \mu\nu$  decays in 2.2 fb<sup>-1</sup> of Run-II data. The mass is determined using the transverse mass, transverse lepton momentum and transverse missing energy distributions, accounting for correlations. This result supersedes AALTONEN 07F. AALTONEN 14D gives more details on the procedures followed by the authors.

<sup>3</sup> ABDALLAH 08A use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events for energies 172 GeV and above. The  $W$  mass was also extracted from the dependence of the  $WW$  cross section close to the production threshold and combined appropriately to obtain the final result. The systematic error includes  $\pm 0.025 \text{ GeV}$  due to final state interactions and  $\pm 0.009 \text{ GeV}$  due to LEP energy uncertainty.

<sup>4</sup> ABBIENDI 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The result quoted here is obtained combining this mass value with the results using  $W^+ W^- \rightarrow \ell\nu_\ell\ell'\nu_{\ell'}$  events in the energy range 183–207 GeV (ABBIENDI 03c) and the dependence of the  $WW$  production cross-section on  $m_{WW}$  at threshold. The systematic error includes  $\pm 0.009 \text{ GeV}$  due to the uncertainty on the LEP beam energy.

<sup>5</sup> ACHARD 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events in the C.M. energy range 189–209 GeV. The result quoted here is obtained combining this mass value with the results obtained from a direct  $W$  mass reconstruction at 172 and 183 GeV and with those from the dependence of the  $WW$  production cross-section on  $m_{WW}$  at 161 and 172 GeV (ACCIARRI 99).

<sup>6</sup> SCHAEEL 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events in the C.M. energy range 183–209 GeV. The result quoted here is obtained combining this mass value with those obtained from the dependence of the  $W$  pair production cross-section on  $m_{WW}$  at 161 and 172 GeV (BARATE 97 and BARATE 97s respectively). The systematic error includes  $\pm 0.009 \text{ GeV}$  due to possible effects of final state interactions in the  $q\bar{q}q\bar{q}$  channel and  $\pm 0.009 \text{ GeV}$  due to the uncertainty on the LEP beam energy.

<sup>7</sup> ABAZOV 02D improve the measurement of the  $W$ -boson mass including  $W \rightarrow e\nu_e$  events in which the electron is close to a boundary of a central electromagnetic calorimeter module. Properly combining the results obtained by fitting  $m_T(W)$ ,  $p_T(e)$ , and  $p_T(\nu)$ , this sample provides a mass value of 80.574 ± 0.405 GeV. The value reported here is a combination of this measurement with all previous D0  $W$ -boson mass measurements.

<sup>8</sup> AFFOLDER 01E fit the transverse mass spectrum of 30115  $W \rightarrow e\nu_e$  events ( $M_{WW} = 80.473 \pm 0.065 \pm 0.092 \text{ GeV}$ ) and of 14740  $W \rightarrow \mu\nu_\mu$  events ( $M_{WW} = 80.465 \pm 0.100 \pm 0.103 \text{ GeV}$ ) obtained in the run IB (1994–95). Combining the electron and muon results, accounting for correlated uncertainties, yields  $M_{WW} = 80.470 \pm 0.089 \text{ GeV}$ . They combine this value with their measurement of ABE 95P reported in run IA (1992–93) to obtain the quoted value.

<sup>9</sup> ABAZOV 12F select 1677k  $W \rightarrow e\nu$  decays in 4.3 fb<sup>-1</sup> of Run-II data. The mass is determined using the transverse mass and transverse lepton momentum distributions, accounting for correlations.

<sup>10</sup> ABAZOV 09AB study the transverse mass, transverse electron momentum, and transverse missing energy in a sample of 0.5 million  $W \rightarrow e\nu$  decays selected in Run-II data. The quoted result combines all three methods, accounting for correlations.

<sup>11</sup> AALTONEN 07F obtain high purity  $W \rightarrow e\nu_e$  and  $W \rightarrow \mu\nu_\mu$  candidate samples totaling 63,964 and 51,128 events respectively. The  $W$  mass value quoted above is derived by simultaneously fitting the transverse mass and the lepton, and neutrino  $p_T$  distributions.

<sup>12</sup> AKTAS 06 fit the  $Q^2$  dependence ( $300 < Q^2 < 30,000 \text{ GeV}^2$ ) of the charged-current differential cross section with a propagator mass. The first error is experimental and the second corresponds to uncertainties due to input parameters and model assumptions.

<sup>13</sup> CHEKANOV 02c fit the  $Q^2$  dependence ( $200 < Q^2 < 60,000 \text{ GeV}^2$ ) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.

<sup>14</sup> BREITWEG 00D fit the  $Q^2$  dependence ( $200 < Q^2 < 22500 \text{ GeV}^2$ ) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.

<sup>15</sup> ALITTI 92B result has two contributions to the systematic error ( $\pm 0.83$ ); one ( $\pm 0.81$ ) cancels in  $m_{WW}/m_Z$  and one ( $\pm 0.17$ ) is noncancelling. These were added in quadrature. We choose the ALITTI 92B value without using the LEP  $m_Z$  value, because we perform our own combined fit.

<sup>16</sup> There are two contributions to the systematic error ( $\pm 0.84$ ): one ( $\pm 0.81$ ) which cancels in  $m_{WW}/m_Z$  and one ( $\pm 0.21$ ) which is non-cancelling. These were added in quadrature.

<sup>17</sup> ABE 89i systematic error dominated by the uncertainty in the absolute energy scale.

<sup>18</sup> ALBAJAR 89 result is from a total sample of 299  $W \rightarrow e\nu$  events.

<sup>19</sup> ALBAJAR 89 result is from a total sample of 67  $W \rightarrow \mu\nu$  events.

<sup>20</sup> ALBAJAR 89 result is from  $W \rightarrow \tau\nu$  events.

## W/Z MASS RATIO

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.88153 ± 0.00017</b>		<sup>1</sup> PDG	16	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.8821 ± 0.0011 ± 0.0008	28323	<sup>2</sup> ABBOTT	98N D0	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
0.88114 ± 0.00154 ± 0.00252	5982	<sup>3</sup> ABBOTT	98P D0	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
0.8813 ± 0.0036 ± 0.0019	156	<sup>4</sup> ALITTI	92B UA2	$E_{\text{cm}}^{p\bar{p}} = 630 \text{ GeV}$

<sup>1</sup> PDG 16 is the PDG average using the world average  $m_W$  and  $m_Z$  values as quoted in this edition of *Review of Particle Physics*. The directly measured values of  $m_W/m_Z$  are not used as their correlation with the Tevatron measured  $m_W$  is unknown.

<sup>2</sup> ABBOTT 98N obtain this from a study of 28323  $W \rightarrow e\nu_e$  and 3294  $Z \rightarrow e^+e^-$  decays. Of this latter sample, 2179 events are used to calibrate the electron energy scale.

<sup>3</sup> ABBOTT 98P obtain this from a study of 5982  $W \rightarrow e\nu_e$  events. The systematic error includes an uncertainty of  $\pm 0.00175$  due to the electron energy scale.

<sup>4</sup> Scale error cancels in this ratio.

## m<sub>Z</sub> – m<sub>W</sub>

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b>10.803 ± 0.015 OUR AVERAGE</b>			
10.803 ± 0.015	<sup>1</sup> PDG	16	
10.4 ± 1.4 ± 0.8	ALBAJAR	89 UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11.3 ± 1.3 ± 0.9	ANSARI	87 UA2	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$

<sup>1</sup> PDG 16 value was obtained using the world average values of  $m_Z$  and  $m_W$  as listed in this publication.

## m<sub>W+</sub> – m<sub>W-</sub>

Test of  $CPT$  invariance.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>–0.19 ± 0.58</b>	1722	ABE	90G CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$

## W WIDTH

The  $W$  width listed here corresponds to the width parameter in a Breit-Wigner distribution with mass-dependent width. To obtain the world average, common systematic uncertainties between experiments are properly taken into account. The LEP-2 average  $W$  width based on published results is  $2.195 \pm 0.083 \text{ GeV}$  [SCHAEEL 13A]. The combined Tevatron data yields an average  $W$  width of  $2.046 \pm 0.049 \text{ GeV}$  [FERMILAB-TM-2460-E].

OUR FIT uses these average LEP and Tevatron width values and combines them assuming no correlations.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.085 ± 0.042 OUR FIT</b>				
2.028 ± 0.072	5272	<sup>1</sup> ABAZOV	09AK D0	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ GeV}$
2.032 ± 0.045 ± 0.057	6055	<sup>2</sup> AALTONEN	08B CDF	$E_{\text{cm}}^{p\bar{p}} = 1.96 \text{ TeV}$
2.404 ± 0.140 ± 0.101	10.3k	<sup>3</sup> ABDALLAH	08A DLPH	$E_{\text{cm}}^{ee} = 183\text{--}209 \text{ GeV}$
1.996 ± 0.096 ± 0.102	10729	<sup>4</sup> ABBIENDI	06 OPAL	$E_{\text{cm}}^{ee} = 170\text{--}209 \text{ GeV}$
2.18 ± 0.11 ± 0.09	9795	<sup>5</sup> ACHARD	06 L3	$E_{\text{cm}}^{ee} = 172\text{--}209 \text{ GeV}$
2.14 ± 0.09 ± 0.06	8717	<sup>6</sup> SCHAEEL	06 ALEP	$E_{\text{cm}}^{ee} = 183\text{--}209 \text{ GeV}$
2.23 $^{+0.15}_{-0.14}$ ± 0.10	294	<sup>7</sup> ABAZOV	02E D0	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
2.05 ± 0.10 ± 0.08	662	<sup>8</sup> AFFOLDER	00M CDF	$E_{\text{cm}}^{p\bar{p}} = 1.8 \text{ TeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.152 ± 0.066	79176	<sup>9</sup> ABBOTT	00B D0	Extracted value
2.064 ± 0.060 ± 0.059		<sup>10</sup> ABE	95W CDF	Extracted value
2.10 $^{+0.14}_{-0.13}$ ± 0.09	3559	<sup>11</sup> ALITTI	92 UA2	Extracted value
2.18 $^{+0.26}_{-0.24}$ ± 0.04		<sup>12</sup> ALBAJAR	91 UA1	Extracted value

<sup>1</sup> ABAZOV 09AK obtain this result fitting the high-end tail (100–200 GeV) of the transverse mass spectrum in  $W \rightarrow e\nu$  decays.

<sup>2</sup> AALTONEN 08B obtain this result fitting the high-end tail (90–200 GeV) of the transverse mass spectrum in semileptonic  $W \rightarrow e\nu_e$  and  $W \rightarrow \mu\nu_\mu$  decays.

<sup>3</sup> ABDALLAH 08A use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The systematic error includes  $\pm 0.065 \text{ GeV}$  due to final state interactions.

<sup>4</sup> ABBIENDI 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The systematic error includes  $\pm 0.003 \text{ GeV}$  due to the uncertainty on the LEP beam energy.

<sup>5</sup> ACHARD 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events in the C.M. energy range 189–209 GeV. The result quoted here is obtained combining this value of the width with the result obtained from a direct  $W$  mass reconstruction at 172 and 183 GeV (ACCIARRI 99).

<sup>6</sup> SCHAEEL 06 use direct reconstruction of the kinematics of  $W^+ W^- \rightarrow q\bar{q}\ell\nu_\ell$  and  $W^+ W^- \rightarrow q\bar{q}q\bar{q}$  events. The systematic error includes  $\pm 0.05 \text{ GeV}$  due to possible effects of final state interactions in the  $q\bar{q}q\bar{q}$  channel and  $\pm 0.01 \text{ GeV}$  due to the uncertainty on the LEP beam energy.



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# Gauge & Higgs Boson Particle Listings

## W

- <sup>7</sup> ABZOV 02e obtain this result fitting the high-end tail (90–200 GeV) of the transverse-mass spectrum in semileptonic  $W \rightarrow e\nu_e$  decays.
- <sup>8</sup> AFFOLDER 00m fit the high transverse mass (100–200 GeV)  $W \rightarrow e\nu_e$  and  $W \rightarrow \mu\nu_\mu$  events to obtain  $\Gamma(W) = 2.04 \pm 0.11(\text{stat}) \pm 0.09(\text{syst})$  GeV. This is combined with the earlier CDF measurement (ABE 95c) to obtain the quoted result.
- <sup>9</sup> ABBOTT 00b measure  $R = 10.43 \pm 0.27$  for the  $W \rightarrow e\nu_e$  decay channel. They use the SM theoretical predictions for  $\sigma(W)/\sigma(Z)$  and  $\Gamma(W \rightarrow e\nu_e)$  and the world average for  $B(Z \rightarrow ee)$ . The value quoted here is obtained combining this result ( $2.169 \pm 0.070$  GeV) with that of ABBOTT 99h.
- <sup>10</sup> ABE 95w measured  $R = 10.90 \pm 0.32 \pm 0.29$ . They use  $m_W = 80.23 \pm 0.18$  GeV,  $\sigma(W)/\sigma(Z) = 3.35 \pm 0.03$ ,  $\Gamma(W \rightarrow e\nu) = 225.9 \pm 0.9$  MeV,  $\Gamma(Z \rightarrow e^+e^-) = 83.98 \pm 0.18$  MeV, and  $\Gamma(Z) = 2.4969 \pm 0.0038$  GeV.
- <sup>11</sup> ALITTI 92 measured  $R = 10.4 \pm 0.7 \pm 0.3$ . The values of  $\sigma(Z)$  and  $\sigma(W)$  come from  $O(\alpha_s^2)$  calculations using  $m_W = 80.14 \pm 0.27$  GeV, and  $m_Z = 91.175 \pm 0.021$  GeV along with the corresponding value of  $\sin^2\theta_W = 0.2274$ . They use  $\sigma(W)/\sigma(Z) = 3.26 \pm 0.07 \pm 0.05$  and  $\Gamma(Z) = 2.487 \pm 0.010$  GeV.
- <sup>12</sup> ALBAJAR 91 measured  $R = 9.5 \pm 1.1$  (stat. + syst.).  $\sigma(W)/\sigma(Z)$  is calculated in QCD at the parton level using  $m_W = 80.18 \pm 0.28$  GeV and  $m_Z = 91.172 \pm 0.031$  GeV along with  $\sin^2\theta_W = 0.2322 \pm 0.0014$ . They use  $\sigma(W)/\sigma(Z) = 3.23 \pm 0.05$  and  $\Gamma(Z) = 2.498 \pm 0.020$  GeV. This measurement is obtained combining both the electron and muon channels.

### W<sup>+</sup> DECAY MODES

$W^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $\ell^+ \nu$	[a] $(10.86 \pm 0.09) \%$	
$\Gamma_2$ $e^+ \nu$	$(10.71 \pm 0.16) \%$	
$\Gamma_3$ $\mu^+ \nu$	$(10.63 \pm 0.15) \%$	
$\Gamma_4$ $\tau^+ \nu$	$(11.38 \pm 0.21) \%$	
$\Gamma_5$ hadrons	$(67.41 \pm 0.27) \%$	
$\Gamma_6$ $\pi^+ \gamma$	$< 7$	$\times 10^{-6}$ 95%
$\Gamma_7$ $D_s^+ \gamma$	$< 1.3$	$\times 10^{-3}$ 95%
$\Gamma_8$ $cX$	$(33.3 \pm 2.6) \%$	
$\Gamma_9$ $c\bar{s}$	$(31 \pm 13 \pm 11) \%$	
$\Gamma_{10}$ invisible	[b] $(1.4 \pm 2.9) \%$	

[a]  $\ell$  indicates each type of lepton ( $e$ ,  $\mu$ , and  $\tau$ ), not sum over them.

[b] This represents the width for the decay of the  $W$  boson into a charged particle with momentum below detectability,  $p < 200$  MeV.

### W PARTIAL WIDTHS

$\Gamma(\text{invisible})$	$\Gamma_{10}$
This represents the width for the decay of the $W$ boson into a charged particle with momentum below detectability, $p < 200$ MeV.	
VALUE (MeV)	DOCUMENT ID TECN COMMENT
$30^{+52}_{-48} \pm 33$	<sup>1</sup> BARATE 99i ALEP $E_{cm}^{ee} = 161+172+183$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
	<sup>2</sup> BARATE 99l ALEP $E_{cm}^{ee} = 161+172+183$ GeV

- <sup>1</sup> BARATE 99i measure this quantity using the dependence of the total cross section  $\sigma_{WW}$  upon a change in the total width. The fit is performed to the  $WW$  measured cross sections at 161, 172, and 183 GeV. This partial width is  $< 139$  MeV at 95%CL.
- <sup>2</sup> BARATE 99l use  $W$ -pair production to search for effectively invisible  $W$  decays, tagging with the decay of the other  $W$  boson to Standard Model particles. The partial width for effectively invisible decay is  $< 27$  MeV at 95%CL.

### W BRANCHING RATIOS

Overall fits are performed to determine the branching ratios of the  $W$  boson. Averages on  $W \rightarrow e\nu$ ,  $W \rightarrow \mu\nu$ , and  $W \rightarrow \tau\nu$ , and their correlations are obtained by combining results from the four LEP experiments properly taking into account the common systematic uncertainties and their correlations [SCHAE 13a]. A first fit determines the three individual leptonic branching ratios  $B(W \rightarrow e\nu)$ ,  $B(W \rightarrow \mu\nu)$ , and  $B(W \rightarrow \tau\nu)$ . This fit has a  $\chi^2 = 6.3$  for 9 degrees of freedom. The correlation coefficients between the branching fractions are 0.14 ( $e-\mu$ ),  $-0.20$  ( $e-\tau$ ),  $-0.12$  ( $\mu-\tau$ ). A second fit assumes lepton universality and determines the leptonic branching ratio  $\text{br } W \rightarrow \ell\nu$  and the hadronic branching ratio is derived as  $B(W \rightarrow \text{hadrons}) = 1-3 \text{ br } W \rightarrow \ell$ . This fit has a  $\chi^2 = 15.4$  for 11 degrees of freedom.

$\Gamma(\ell^+ \nu)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
$\ell$ indicates average over $e$ , $\mu$ , and $\tau$ modes, not sum over modes.	
VALUE (units $10^{-2}$ )	EVTS DOCUMENT ID TECN COMMENT
<b><math>10.86 \pm 0.09</math> OUR FIT</b>	
$10.86 \pm 0.12 \pm 0.08$	16438 ABBIENDI 07A OPAL $E_{cm}^{ee} = 161-209$ GeV
$10.85 \pm 0.14 \pm 0.08$	13600 ABDALLAH 04G DLPH $E_{cm}^{ee} = 161-209$ GeV
$10.83 \pm 0.14 \pm 0.10$	11246 ACHARD 04J L3 $E_{cm}^{ee} = 161-209$ GeV
$10.96 \pm 0.12 \pm 0.05$	16116 SCHAE 04A ALEP $E_{cm}^{ee} = 183-209$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	

$11.02 \pm 0.52$	11858 <sup>1</sup> ABBOTT 99H D0 $E_{cm}^{p\bar{p}} = 1.8$ TeV
$10.4 \pm 0.8$	3642 <sup>2</sup> ABE 92i CDF $E_{cm}^{p\bar{p}} = 1.8$ TeV

- <sup>1</sup> ABBOTT 99H measure  $R \equiv [\sigma_W B(W \rightarrow \ell\nu_\ell)]/[\sigma_Z B(Z \rightarrow \ell\ell)] = 10.90 \pm 0.52$  combining electron and muon channels. They use  $M_W = 80.39 \pm 0.06$  GeV and the SM theoretical predictions for  $\sigma(W)/\sigma(Z)$  and  $B(Z \rightarrow \ell\ell)$ .
- <sup>2</sup>  $1216 \pm 38^{+27}_{-31}$   $W \rightarrow \mu\nu$  events from ABE 92i and  $2426 W \rightarrow e\nu$  events of ABE 91c. ABE 92i give the inverse quantity as  $9.6 \pm 0.7$  and we have inverted.

$\Gamma(e^+ \nu)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE (units $10^{-2}$ )	EVTS DOCUMENT ID TECN COMMENT
<b><math>10.71 \pm 0.16</math> OUR FIT</b>	
$10.71 \pm 0.25 \pm 0.11$	2374 ABBIENDI 07A OPAL $E_{cm}^{ee} = 161-209$ GeV
$10.55 \pm 0.31 \pm 0.14$	1804 ABDALLAH 04G DLPH $E_{cm}^{ee} = 161-209$ GeV
$10.78 \pm 0.29 \pm 0.13$	1576 ACHARD 04J L3 $E_{cm}^{ee} = 161-209$ GeV
$10.78 \pm 0.27 \pm 0.10$	2142 SCHAE 04A ALEP $E_{cm}^{ee} = 183-209$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$10.61 \pm 0.28$	<sup>1</sup> ABZOV 04d TEVA $E_{cm}^{p\bar{p}} = 1.8$ TeV

- <sup>1</sup> ABZOV 04d take into account all correlations to properly combine the CDF (ABE 95w) and DØ (ABBOTT 00b) measurements of the ratio  $R$  in the electron channel. The ratio  $R$  is defined as  $[\sigma_W \cdot B(W \rightarrow e\nu_e)] / [\sigma_Z \cdot B(Z \rightarrow ee)]$ . The combination gives  $R^{\text{TeVatron}} = 10.59 \pm 0.23$ .  $\sigma_W / \sigma_Z$  is calculated at next-to-next-to-leading order ( $3.360 \pm 0.051$ ). The branching fraction  $B(Z \rightarrow ee)$  is taken from this Review as ( $3.363 \pm 0.004$ )%.

$\Gamma(\mu^+ \nu)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE (units $10^{-2}$ )	EVTS DOCUMENT ID TECN COMMENT
<b><math>10.63 \pm 0.15</math> OUR FIT</b>	
$10.78 \pm 0.24 \pm 0.10$	2397 ABBIENDI 07A OPAL $E_{cm}^{ee} = 161-209$ GeV
$10.65 \pm 0.26 \pm 0.08$	1998 ABDALLAH 04G DLPH $E_{cm}^{ee} = 161-209$ GeV
$10.03 \pm 0.29 \pm 0.12$	1423 ACHARD 04J L3 $E_{cm}^{ee} = 161-209$ GeV
$10.87 \pm 0.25 \pm 0.08$	2216 SCHAE 04A ALEP $E_{cm}^{ee} = 183-209$ GeV

$\Gamma(\tau^+ \nu)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
VALUE (units $10^{-2}$ )	EVTS DOCUMENT ID TECN COMMENT
<b><math>11.38 \pm 0.21</math> OUR FIT</b>	
$11.14 \pm 0.31 \pm 0.17$	2177 ABBIENDI 07A OPAL $E_{cm}^{ee} = 161-209$ GeV
$11.46 \pm 0.39 \pm 0.19$	2034 ABDALLAH 04G DLPH $E_{cm}^{ee} = 161-209$ GeV
$11.89 \pm 0.40 \pm 0.20$	1375 ACHARD 04J L3 $E_{cm}^{ee} = 161-209$ GeV
$11.25 \pm 0.32 \pm 0.20$	2070 SCHAE 04A ALEP $E_{cm}^{ee} = 183-209$ GeV

$\Gamma(\text{hadrons})/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
OUR FIT value is obtained by a fit to the lepton branching ratio data assuming lepton universality.	

VALUE (units $10^{-2}$ )	EVTS DOCUMENT ID TECN COMMENT
<b><math>67.41 \pm 0.27</math> OUR FIT</b>	
$67.41 \pm 0.37 \pm 0.23$	16438 ABBIENDI 07A OPAL $E_{cm}^{ee} = 161-209$ GeV
$67.45 \pm 0.41 \pm 0.24$	13600 ABDALLAH 04G DLPH $E_{cm}^{ee} = 161-209$ GeV
$67.50 \pm 0.42 \pm 0.30$	11246 ACHARD 04J L3 $E_{cm}^{ee} = 161-209$ GeV
$67.13 \pm 0.37 \pm 0.15$	16116 SCHAE 04A ALEP $E_{cm}^{ee} = 183-209$ GeV

$\Gamma(\mu^+ \nu)/\Gamma(e^+ \nu)$	$\Gamma_3/\Gamma_2$
VALUE	EVTS DOCUMENT ID TECN COMMENT
<b><math>0.991 \pm 0.018</math> OUR AVERAGE</b>	
$0.993 \pm 0.019$	SCHAE 13A LEP $E_{cm}^{ee} = 130-209$ GeV
$0.89 \pm 0.10$	13k <sup>1</sup> ABACHI 95D D0 $E_{cm}^{p\bar{p}} = 1.8$ TeV
$1.02 \pm 0.08$	1216 <sup>2</sup> ABE 92i CDF $E_{cm}^{p\bar{p}} = 1.8$ TeV
$1.00 \pm 0.14 \pm 0.08$	67 ALBAJAR 89 UA1 $E_{cm}^{p\bar{p}} = 546,630$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$1.24 \pm 0.6 \pm 0.4$	14 ARNISON 84D UA1 Repl. by ALBAJAR 89

- <sup>1</sup> ABACHI 95D obtain this result from the measured  $\sigma_{WB}(W \rightarrow \mu\nu) = 2.09 \pm 0.23 \pm 0.11$  nb and  $\sigma_{WB}(W \rightarrow e\nu) = 2.36 \pm 0.07 \pm 0.13$  nb in which the first error is the combined statistical and systematic uncertainty, the second reflects the uncertainty in the luminosity.
- <sup>2</sup> ABE 92i obtain  $\sigma_{WB}(W \rightarrow \mu\nu) = 2.21 \pm 0.07 \pm 0.21$  and combine with ABE 91c  $\sigma_W B(W \rightarrow e\nu)$  to give a ratio of the couplings from which we derive this measurement.

$\Gamma(\tau^+ \nu)/\Gamma(e^+ \nu)$	$\Gamma_4/\Gamma_2$
VALUE	EVTS DOCUMENT ID TECN COMMENT
<b><math>1.043 \pm 0.024</math> OUR AVERAGE</b>	
$1.063 \pm 0.027$	SCHAE 13A LEP $E_{cm}^{ee} = 130-209$ GeV
$0.961 \pm 0.061$	980 <sup>1</sup> ABBOTT 00D D0 $E_{cm}^{p\bar{p}} = 1.8$ TeV
$0.94 \pm 0.14$	179 <sup>2</sup> ABE 92e CDF $E_{cm}^{p\bar{p}} = 1.8$ TeV
$1.04 \pm 0.08 \pm 0.08$	754 <sup>3</sup> ALITTI 92F UA1 $E_{cm}^{p\bar{p}} = 630$ GeV
$1.02 \pm 0.20 \pm 0.12$	32 ALBAJAR 89 UA1 $E_{cm}^{p\bar{p}} = 546,630$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.995 \pm 0.112 \pm 0.083$	198 ALITTI 91c UA2 Repl. by ALITTI 92F
$1.02 \pm 0.20 \pm 0.10$	32 ALBAJAR 87 UA1 Repl. by ALBAJAR 89

# Gauge & Higgs Boson Particle Listings

## W

<sup>1</sup> ABBOTT 00D measure  $\sigma_{W \times B}(W \rightarrow \tau \nu_\tau) = 2.22 \pm 0.09 \pm 0.10 \pm 0.10$  nb. Using the ABBOTT 00B result  $\sigma_{W \times B}(W \rightarrow e \nu_e) = 2.31 \pm 0.01 \pm 0.05 \pm 0.10$  nb, they quote the ratio of the couplings from which we derive this measurement.

<sup>2</sup> ABE 92E use two procedures for selecting  $W \rightarrow \tau \nu_\tau$  events. The missing  $E_T$  trigger leads to  $132 \pm 14 \pm 8$  events and the  $\tau$  trigger to  $47 \pm 9 \pm 4$  events. Proper statistical and systematic correlations are taken into account to arrive at  $\sigma_B(W \rightarrow \tau \nu) = 2.05 \pm 0.27$  nb. Combined with ABE 91C result on  $\sigma_B(W \rightarrow e \nu)$ , ABE 92E quote a ratio of the couplings from which we derive this measurement.

<sup>3</sup> This measurement is derived by us from the ratio of the couplings of ALITTI 92F.

$\Gamma(\tau^+ \nu)/\Gamma(\mu^+ \nu)$		DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma_3$
VALUE					
<b>1.070 ± 0.026</b>		SCHAE	13A	LEP	$E_{cm}^e = 130-209$ GeV

$\Gamma(\pi^+ \gamma)/\Gamma(e^+ \nu)$		DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma_2$
VALUE	CL%				
<b>&lt; 6.4 × 10<sup>-5</sup></b>	95	AALTONEN	12W	CDF	$E_{cm}^{pp} = 1.96$ TeV
< 7 × 10 <sup>-4</sup>	95	ABE	98H	CDF	$E_{cm}^{pp} = 1.8$ TeV
< 4.9 × 10 <sup>-3</sup>	95	<sup>1</sup> ALITTI	92D	UA2	$E_{cm}^{pp} = 630$ GeV
< 58 × 10 <sup>-3</sup>	95	<sup>2</sup> ALBAJAR	90	UA1	$E_{cm}^{pp} = 546, 630$ GeV

<sup>1</sup> ALITTI 92D limit is  $3.8 \times 10^{-3}$  at 90%CL.

<sup>2</sup> ALBAJAR 90 obtain < 0.048 at 90%CL.

$\Gamma(D_s^+ \gamma)/\Gamma(e^+ \nu)$		DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma_2$
VALUE	CL%				
<b>&lt; 1.2 × 10<sup>-2</sup></b>	95	ABE	98P	CDF	$E_{cm}^{pp} = 1.8$ TeV

$\Gamma(cX)/\Gamma(\text{hadrons})$		DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma_5$
VALUE	EVTS				
<b>0.49 ± 0.04 OUR AVERAGE</b>					
0.481 ± 0.042 ± 0.032	3005	<sup>1</sup> ABBIENDI	00V	OPAL	$E_{cm}^e = 183 + 189$ GeV
0.51 ± 0.05 ± 0.03	746	<sup>2</sup> BARATE	99M	ALEP	$E_{cm}^e = 172 + 183$ GeV

<sup>1</sup> ABBIENDI 00v tag  $W \rightarrow cX$  decays using measured jet properties, lifetime information, and leptons produced in charm decays. From this result, and using the additional measurements of  $\Gamma(W)$  and  $B(W \rightarrow \text{hadrons})$ ,  $|V_{cs}|$  is determined to be  $0.969 \pm 0.045 \pm 0.036$ .

<sup>2</sup> BARATE 99m tag c jets using a neural network algorithm. From this measurement  $|V_{cs}|$  is determined to be  $1.00 \pm 0.11 \pm 0.07$ .

$R_{cs} = \Gamma(c\bar{s})/\Gamma(\text{hadrons})$		DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma_5$
VALUE					
<b>0.46 ± 0.18 ± 0.14 ± 0.07</b>		<sup>1</sup> ABREU	98N	DLPH	$E_{cm}^e = 161+172$ GeV

<sup>1</sup> ABREU 98N tag c and s jets by identifying a charged kaon as the highest momentum particle in a hadronic jet. They also use a lifetime tag to independently identify a c jet, based on the impact parameter distribution of charged particles in a jet. From this measurement  $|V_{cs}|$  is determined to be  $0.94^{+0.32}_{-0.26} \pm 0.13$ .

### AVERAGE PARTICLE MULTIPLICITIES IN HADRONIC W DECAY

Summed over particle and antiparticle, when appropriate.

$\langle N_{\pi^\pm} \rangle$		DOCUMENT ID	TECN	COMMENT
VALUE				
<b>15.70 ± 0.35</b>		<sup>1</sup> ABREU,P	00F	DLPH $E_{cm}^e = 189$ GeV

<sup>1</sup> ABREU,P 00F measure  $\langle N_{\pi^\pm} \rangle = 31.65 \pm 0.48 \pm 0.76$  and  $15.51 \pm 0.38 \pm 0.40$  in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

$\langle N_{K^\pm} \rangle$		DOCUMENT ID	TECN	COMMENT
VALUE				
<b>2.20 ± 0.19</b>		<sup>1</sup> ABREU,P	00F	DLPH $E_{cm}^e = 189$ GeV

<sup>1</sup> ABREU,P 00F measure  $\langle N_{K^\pm} \rangle = 4.38 \pm 0.42 \pm 0.12$  and  $2.23 \pm 0.32 \pm 0.17$  in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

$\langle N_p \rangle$		DOCUMENT ID	TECN	COMMENT
VALUE				
<b>0.92 ± 0.14</b>		<sup>1</sup> ABREU,P	00F	DLPH $E_{cm}^e = 189$ GeV

<sup>1</sup> ABREU,P 00F measure  $\langle N_p \rangle = 1.82 \pm 0.29 \pm 0.16$  and  $0.94 \pm 0.23 \pm 0.06$  in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

$\langle N_{\text{charged}} \rangle$		DOCUMENT ID	TECN	COMMENT
VALUE				
<b>19.39 ± 0.08 OUR AVERAGE</b>				
19.38 ± 0.05 ± 0.08		<sup>1</sup> ABBIENDI	06A	OPAL $E_{cm}^e = 189-209$ GeV
19.44 ± 0.17		<sup>2</sup> ABREU,P	00F	DLPH $E_{cm}^e = 183+189$ GeV
19.3 ± 0.3 ± 0.3		<sup>3</sup> ABBIENDI	99N	OPAL $E_{cm}^e = 183$ GeV
19.23 ± 0.74		<sup>4</sup> ABREU	98C	DLPH $E_{cm}^e = 172$ GeV

<sup>1</sup> ABBIENDI 06A measure  $\langle N_{\text{charged}} \rangle = 38.74 \pm 0.12 \pm 0.26$  when both  $W$  bosons decay hadronically and  $\langle N_{\text{charged}} \rangle = 19.39 \pm 0.11 \pm 0.09$  when one  $W$  boson decays semileptonically. The value quoted here is obtained under the assumption that there is no color reconnection between  $W$  bosons; the value is a weighted average taking into account correlations in the systematic uncertainties.

<sup>2</sup> ABREU,P 00F measure  $\langle N_{\text{charged}} \rangle = 39.12 \pm 0.33 \pm 0.36$  and  $38.11 \pm 0.57 \pm 0.44$  in the fully hadronic final states at 189 and 183 GeV respectively, and  $\langle N_{\text{charged}} \rangle = 19.49 \pm 0.31 \pm 0.27$  and  $19.78 \pm 0.49 \pm 0.43$  in the semileptonic final states. The value quoted is a weighted average without assuming any correlations.

<sup>3</sup> ABBIENDI 99N use the final states  $W^+ W^- \rightarrow q\bar{q} \ell \bar{\nu}_\ell$  to derive this value.

<sup>4</sup> ABREU 98C combine results from both the fully hadronic as well semileptonic  $WW$  final states after demonstrating that the  $W$  decay charged multiplicity is independent of the topology within errors.

### TRIPLE GAUGE COUPLINGS (TGC'S) EXTRACTION OF TRIPLE GAUGE COUPLINGS (TGCS)

Revised August 2015 by M.W. Grünewald (U. College Dublin) and A. Gurtu (Formerly Tata Inst.).

Fourteen independent couplings, seven each for  $ZWW$  and  $\gamma WW$ , completely describe the  $VWW$  vertices within the most general framework of the electroweak Standard Model (SM) consistent with Lorentz invariance and U(1) gauge invariance. Of each of the seven TGCs, three conserve  $C$  and  $P$  individually, three violate  $CP$ , and one violates  $C$  and  $P$  individually while conserving  $CP$ . Assumption of  $C$  and  $P$  conservation and electromagnetic gauge invariance reduces the number of independent  $VWW$  couplings to five: one common set [1,2] is  $(\kappa_\gamma, \kappa_Z, \lambda_\gamma, \lambda_Z, g_1^Z)$ , where  $\kappa_\gamma = \kappa_Z = g_1^Z = 1$  and  $\lambda_\gamma = \lambda_Z = 0$  in the Standard Model at tree level. The parameters  $\kappa_Z$  and  $\lambda_Z$  are related to the other three due to constraints of gauge invariance as follows:  $\kappa_Z = g_1^Z - (\kappa_\gamma - 1) \tan^2 \theta_W$  and  $\lambda_Z = \lambda_\gamma$ , where  $\theta_W$  is the weak mixing angle. The  $W$  magnetic dipole moment,  $\mu_W$ , and the  $W$  electric quadrupole moment,  $q_W$ , are expressed as  $\mu_W = e(1 + \kappa_\gamma + \lambda_\gamma)/2M_W$  and  $q_W = -e(\kappa_\gamma - \lambda_\gamma)/M_W^2$ .

Precision measurements of suitable observables at LEP1 has already led to an exploration of much of the TGC parameter space. At LEP2, the  $VWW$  coupling arises in  $W$ -pair production via  $s$ -channel exchange, or in single  $W$  production via the radiation of a virtual photon off the incident  $e^+$  or  $e^-$ . At the Tevatron and the LHC, hard-photon bremsstrahlung off a produced  $W$  or  $Z$  signals the presence of a triple-gauge vertex. In order to extract the value of one TGC, the others are generally kept fixed to their SM values. While most analyses use the above gauge constraints in the extraction of TGCs, one analysis of  $W$ -pair events also determines the real and imaginary parts of all 14 couplings using unconstrained single-parameter fits [3]. The results are consistent. Some experiments have determined limits on the couplings under various non-LEP scenarios and assuming different values of the form factor  $\Lambda$ , where the coupling parameters are scaled by  $1/(1 + s/\Lambda^2)^2$ . For practical reasons it is not possible to quote all such determinations in the listings. For that the individual papers may be consulted.

### References

1. K. Hagiwara *et al.*, Nucl. Phys. **B282**, 253 (1987).
2. G. Gounaris *et al.*, CERN 96-01 p. 525.
3. S. Schael *et al.* (ALEPH Collab.), Phys. Lett. **B614**, 7 (2005).

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Gauge & Higgs Boson Particle Listings  
*W*

**$g_1^Z$**   
OUR FIT below is taken from [SCHAE1 13A].

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.984<math>^{+0.018}_{-0.020}</math> OUR FIT</b>				
0.975 $^{+0.033}_{-0.030}$	7872	<sup>1</sup> ABDALLAH	10 DLPH	$E_{\text{cm}}^{\text{ee}} = 189\text{--}209$ GeV
1.001 $\pm 0.027 \pm 0.013$	9310	<sup>2</sup> SCHAE1	05A ALEP	$E_{\text{cm}}^{\text{ee}} = 183\text{--}209$ GeV
0.987 $^{+0.034}_{-0.033}$	9800	<sup>3</sup> ABBIENDI	04D OPAL	$E_{\text{cm}}^{\text{ee}} = 183\text{--}209$ GeV
0.966 $^{+0.034}_{-0.032} \pm 0.015$	8325	<sup>4</sup> ACHARD	04D L3	$E_{\text{cm}}^{\text{ee}} = 161\text{--}209$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>5</sup> AAD	14Y ATLS	$E_{\text{cm}}^{\text{pp}} = 8$ TeV
		<sup>6</sup> AAD	13AL ATLS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
		<sup>7</sup> CHATRCHYAN13BF	CMS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
		<sup>8</sup> AAD	12CD ATLS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
		<sup>9</sup> AALTONEN	12AC CDF	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
		<sup>10</sup> ABAZOV	12AG D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
	34	<sup>11</sup> ABAZOV	11 D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
	334	<sup>12</sup> AALTONEN	10K CDF	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
1.04 $\pm 0.09$		<sup>13</sup> ABAZOV	09AD D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
		<sup>14</sup> ABAZOV	09AJ D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
1.07 $^{+0.08}_{-0.12}$	1880	<sup>15</sup> ABDALLAH	08C DLPH	Superseded by ABDAL-LAH 10
	13	<sup>16</sup> ABAZOV	07Z D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
	2.3	<sup>17</sup> ABAZOV	05S D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
0.98 $\pm 0.07 \pm 0.01$	2114	<sup>18</sup> ABREU	01I DLPH	$E_{\text{cm}}^{\text{ee}} = 183\text{--}189$ GeV
	331	<sup>19</sup> ABBOTT	99I D0	$E_{\text{cm}}^{\text{pp}} = 1.8$ TeV

<sup>1</sup> ABDALLAH 10 use data on the final states  $e^+e^- \rightarrow jj\ell\nu, jjjj, jjX, \ell X$ , at center-of-mass energies between 189–209 GeV at LEP2, where  $j = \text{jet}$ ,  $\ell = \text{lepton}$ , and  $X$  represents missing momentum. The fit is carried out keeping all other parameters fixed at their SM values.

<sup>2</sup> SCHAE1 05A study single-photon, single- $W$ , and  $WW$ -pair production from 183 to 209 GeV. The result quoted here is derived from the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

<sup>3</sup> ABBIENDI 04D combine results from  $W^+W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $0.923 < g_1^Z < 1.054$ .

<sup>4</sup> ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained from the  $WW$ -pair production sample including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

<sup>5</sup> AAD 14Y determine the electroweak  $Z$ -dijet cross section in 8 TeV  $pp$  collisions.  $Z \rightarrow ee$  and  $Z \rightarrow \mu\mu$  decays are selected with the di-lepton  $p_T > 20$  GeV and mass in the 81–101 GeV range. Minimum two jets are required with  $p_T > 55$  and 45 GeV and no additional jets with  $p_T > 25$  GeV in the rapidity interval between them. The normalized  $p_T$  balance between the  $Z$  and the two jets is required to be  $< 0.15$ . This leads to a selection of 900 events with dijet mass  $> 1$  TeV. The number of signal and background events expected is 261 and 592 respectively. A Poisson likelihood method is used on an event by event basis to obtain the 95% CL limit  $0.5 < g_1^Z < 1.26$  for a form factor value  $\Lambda = \infty$ .

<sup>6</sup> AAD 13AL study  $WW$  production in  $pp$  collisions and select 1325  $WW$  candidates in decay modes with electrons or muons with an expected background of  $369 \pm 61$  events. Assuming the LEP formulation and setting the form-factor  $\Lambda = \text{infinity}$ , a fit to the transverse momentum distribution of the leading charged lepton, leads to a 95% C.L. range of  $0.961 < g_1^Z < 1.052$ . Supersedes AAD 12AC.

<sup>7</sup> CHATRCHYAN 13BF determine the  $W^+W^-$  production cross section using unlike sign di-lepton ( $e$  or  $\mu$ ) events with high  $p_T$ . The leptons have  $p_T > 20$  GeV/c and are isolated. 1134 candidate events are observed with an expected SM background of  $247 \pm 34$ . The  $p_T$  distribution of the leading lepton is fitted to obtain 95% C.L. limits of  $0.905 \leq g_1^Z \leq 1.095$ .

<sup>8</sup> AAD 12CD study  $WZ$  production in  $pp$  collisions and select 317  $WZ$  candidates in three  $\ell\nu$  decay modes with an expected background of  $68.0 \pm 10.0$  events. The resulting 95% C.L. range is:  $0.943 < g_1^Z < 1.093$ . Supersedes AAD 12V.

<sup>9</sup> AALTONEN 12AC study  $WZ$  production in  $p\bar{p}$  collisions and select 63  $WZ$  candidates in three  $\ell\nu$  decay modes with an expected background of  $7.9 \pm 1.0$  events. Based on the cross section and shape of the  $Z$  transverse momentum spectrum, the following 95% C.L. range is reported:  $0.92 < g_1^Z < 1.20$  for a form factor of  $\Lambda = 2$  TeV.

<sup>10</sup> ABAZOV 12AG combine new results with already published results on  $W\gamma, WW$  and  $WZ$  production in order to determine the couplings with increased precision, superseding ABAZOV 08R, ABAZOV 11AC, ABAZOV 09AJ, ABAZOV 09AD. The 68% C.L. result for a formfactor cutoff of  $\Lambda = 2$  TeV is  $g_1^Z = 1.022^{+0.032}_{-0.030}$ .

<sup>11</sup> ABAZOV 11 study the  $p\bar{p} \rightarrow 3\nu$  process arising in  $WZ$  production. They observe 34  $WZ$  candidates with an estimated background of 6 events. An analysis of the  $p_T$  spectrum of the  $Z$  boson leads to a 95% C.L. limit of  $0.944 < g_1^Z < 1.154$ , for a form factor  $\Lambda = 2$  TeV.

<sup>12</sup> AALTONEN 10K study  $p\bar{p} \rightarrow W^+W^-$  with  $W \rightarrow e/\mu\nu$ . The  $p_T$  of the leading (second) lepton is required to be  $> 20$  (10) GeV. The final number of events selected is 654 of which  $320 \pm 47$  are estimated to be background. The 95% C.L. interval is  $0.76 < g_1^Z < 1.34$  for  $\Lambda = 1.5$  TeV and  $0.78 < g_1^Z < 1.30$  for  $\Lambda = 2$  TeV.

<sup>13</sup> ABAZOV 09AD study the  $p\bar{p} \rightarrow \ell\nu 2\text{jet}$  process arising in  $WW$  and  $WZ$  production. They select 12,473 (14,392) events in the electron (muon) channel with an expected

di-boson signal of  $436$  (527) events. The results on the anomalous couplings are derived from an analysis of the  $p_T$  spectrum of the 2-jet system and quoted at 68% C.L. and for a form factor of 2 TeV. This measurement is not used for obtaining the mean as it is for a specific form factor. The 95% confidence interval is  $0.88 < g_1^Z < 1.20$ .

<sup>14</sup> ABAZOV 09AJ study the  $p\bar{p} \rightarrow 2\ell 2\nu$  process arising in  $WW$  production. They select 100 events with an expected  $WW$  signal of 65 events. An analysis of the  $p_T$  spectrum of the two charged leptons leads to 95% C.L. limits of  $0.86 < g_1^Z < 1.3$ , for a form factor  $\Lambda = 2$  TeV.

<sup>15</sup> ABDALLAH 08C determine this triple gauge coupling from the measurement of the spin density matrix elements in  $e^+e^- \rightarrow W^+W^- \rightarrow (qq)(\ell\nu)$ , where  $\ell = e$  or  $\mu$ . Values of all other couplings are fixed to their standard model values.

<sup>16</sup> ABAZOV 07Z set limits on anomalous TGCs using the measured cross section and  $p_T(Z)$  distribution in  $WZ$  production with both the  $W$  and the  $Z$  decaying leptonically into electrons and muons. Setting the other couplings to their standard model values, the 95% C.L. limit for a form factor scale  $\Lambda = 2$  TeV is  $0.86 < g_1^Z < 1.35$ .

<sup>17</sup> ABAZOV 05S study  $p\bar{p} \rightarrow WZ$  production with a subsequent trilepton decay to  $\ell\nu\ell'\bar{\ell}'$  ( $\ell$  and  $\ell' = e$  or  $\mu$ ). Three events (estimated background  $0.71 \pm 0.08$  events) with  $WZ$  decay characteristics are observed from which they derive limits on the anomalous  $WWZ$  couplings. The 95% CL limit for a form factor scale  $\Lambda = 1.5$  TeV is  $0.51 < g_1^Z < 1.66$ , fixing  $\lambda_Z$  and  $\kappa_Z$  to their Standard Model values.

<sup>18</sup> ABREU 01I combine results from  $e^+e^-$  interactions at 189 GeV leading to  $W^+W^-$  and  $W\nu_e$  final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is  $0.84 < g_1^Z < 1.13$ .

<sup>19</sup> ABBOTT 99I perform a simultaneous fit to the  $W\gamma, WW \rightarrow \text{dilepton}, WW/WZ \rightarrow e\nu jj, WW/WZ \rightarrow \mu\nu jj$ , and  $WZ \rightarrow \text{trilepton}$  data samples. For  $\Lambda = 2.0$  TeV, the 95%CL limits are  $0.63 < g_1^Z < 1.57$ , fixing  $\lambda_Z$  and  $\kappa_Z$  to their Standard Model values, and assuming Standard Model values for the  $WW\gamma$  couplings.

**$\kappa_\gamma$**   
OUR FIT below is taken from [SCHAE1 13A].

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.982<math>\pm 0.042</math> OUR FIT</b>				
1.024 $^{+0.077}_{-0.081}$	7872	<sup>1</sup> ABDALLAH	10 DLPH	$E_{\text{cm}}^{\text{ee}} = 189\text{--}209$ GeV
0.971 $\pm 0.055 \pm 0.030$	10689	<sup>2</sup> SCHAE1	05A ALEP	$E_{\text{cm}}^{\text{ee}} = 183\text{--}209$ GeV
0.88 $^{+0.09}_{-0.08}$	9800	<sup>3</sup> ABBIENDI	04D OPAL	$E_{\text{cm}}^{\text{ee}} = 183\text{--}209$ GeV
1.013 $^{+0.067}_{-0.064} \pm 0.026$	10575	<sup>4</sup> ACHARD	04D L3	$E_{\text{cm}}^{\text{ee}} = 161\text{--}209$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>5</sup> CHATRCHYAN14AB	CMS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
		<sup>6</sup> AAD	13AN ATLS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
		<sup>7</sup> CHATRCHYAN13BF	CMS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
		<sup>8</sup> ABAZOV	12AG D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
		<sup>9</sup> ABAZOV	11AC D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
		<sup>10</sup> CHATRCHYAN11M	CMS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV
	334	<sup>11</sup> AALTONEN	10K CDF	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
	53	<sup>12</sup> AARON	09B H1	$E_{\text{cm}}^{\text{pp}} = 0.3$ TeV
1.07 $^{+0.26}_{-0.29}$		<sup>13</sup> ABAZOV	09AD D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
		<sup>14</sup> ABAZOV	09AJ D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
		<sup>15</sup> ABAZOV	08R D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
0.68 $^{+0.17}_{-0.15}$	1880	<sup>16</sup> ABDALLAH	08C DLPH	Superseded by ABDAL-LAH 10
	1617	<sup>17</sup> AALTONEN	07L CDF	$E_{\text{cm}}^{\text{pp}} = 1.96$ GeV
	17	<sup>18</sup> ABAZOV	06H D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
	141	<sup>19</sup> ABAZOV	05J D0	$E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
1.25 $^{+0.21}_{-0.20} \pm 0.06$	2298	<sup>20</sup> ABREU	01I DLPH	$E_{\text{cm}}^{\text{ee}} = 183\text{--}189$ GeV
		<sup>21</sup> BREITWEG	00 ZEUS	$e^+p \rightarrow e^+W^\pm X, \sqrt{s} \approx 300$ GeV
0.92 $\pm 0.34$	331	<sup>22</sup> ABBOTT	99I D0	$E_{\text{cm}}^{\text{pp}} = 1.8$ TeV

<sup>1</sup> ABDALLAH 10 use data on the final states  $e^+e^- \rightarrow jj\ell\nu, jjjj, jjX, \ell X$ , at center-of-mass energies between 189–209 GeV at LEP2, where  $j = \text{jet}$ ,  $\ell = \text{lepton}$ , and  $X$  represents missing momentum. The fit is carried out keeping all other parameters fixed at their SM values.

<sup>2</sup> SCHAE1 05A study single-photon, single- $W$ , and  $WW$ -pair production from 183 to 209 GeV. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

<sup>3</sup> ABBIENDI 04D combine results from  $W^+W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $0.73 < \kappa_\gamma < 1.07$ .

<sup>4</sup> ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

<sup>5</sup> CHATRCHYAN 14AB measure  $W\gamma$  production cross section for  $p_T^\gamma > 15$  GeV and  $R(\ell\gamma) > 0.7$ , which is the separation between the  $\gamma$  and the final state charged lepton ( $e$  or  $\mu$ ) in the azimuthal angle-pseudorapidity ( $\phi - \eta$ ) plane. After background subtraction the number of  $e\nu\gamma$  and  $\mu\nu\gamma$  events is determined to be  $3200 \pm 325$  and  $4970 \pm 543$  respectively, compatible with expectations from the SM. This leads to a 95% CL limit of  $0.62 < \kappa_\gamma < 1.29$ , assuming other parameters have SM values.

# Gauge & Higgs Boson Particle Listings

## W

- <sup>6</sup> AAD 13AN study  $W\gamma$  production in  $pp$  collisions. In events with no additional jet, 4449 (6578)  $W$  decays to electron (muon) are selected, with an expected background of  $1662 \pm 262$  (2538  $\pm$  362) events. Analysing the photon  $p_T$  spectrum above 100 GeV yields a 95% C.L. limit of  $0.59 < \kappa_\gamma < 1.46$ . Supersedes AAD 12Bx.
- <sup>7</sup> CHATRCHYAN 13BF determine the  $W^+W^-$  production cross section using unlike sign di-lepton ( $e$  or  $\mu$ ) events with high  $p_T$ . The leptons have  $p_T > 20$  GeV/c and are isolated. 1134 candidate events are observed with an expected SM background of 247  $\pm$  34. The  $p_T$  distribution of the leading lepton is fitted to obtain 95% C.L. limits of  $0.79 \leq \kappa_\gamma \leq 1.22$ .
- <sup>8</sup> ABAZOV 12AG combine new results with already published results on  $W\gamma$ ,  $WW$  and  $WZ$  production in order to determine the couplings with increased precision, superseding ABAZOV 08R, ABAZOV 11AC, ABAZOV 09AJ, ABAZOV 09AD. The 68% C.L. result for a formfactor cutoff of  $\Lambda = 2$  TeV is  $\kappa_\gamma = 1.048^{+0.106}_{-0.105}$ .
- <sup>9</sup> ABAZOV 11AC study  $W\gamma$  production in  $p\bar{p}$  collisions at 1.96 TeV, with the  $W$  decay products containing an electron or a muon. They select 196 (363) events in the electron (muon) mode, with a SM expectation of 190 (372) events. A likelihood fit to the photon  $E_T$  spectrum above 15 GeV yields at 95% C.L. the result:  $0.6 < \kappa_\gamma < 1.4$  for a formfactor  $\Lambda = 2$  TeV.
- <sup>10</sup> CHATRCHYAN 11M study  $W\gamma$  production in  $pp$  collisions at  $\sqrt{s} = 7$  TeV using 36 pb<sup>-1</sup>  $pp$  data with the  $W$  decaying to electron and muon. The total cross section is measured for photon transverse energy  $E_T^\gamma > 10$  GeV and spatial separation from charged leptons in the plane of pseudo rapidity and azimuthal angle  $\Delta R(\ell, \gamma) > 0.7$ . The number of candidate (background) events is 452 (228  $\pm$  21) for the electron channel and 520 (277  $\pm$  25) for the muon channel. Setting other couplings to their standard model value, they derive a 95% CL limit of  $-0.11 < \kappa_\gamma < 0.204$ .
- <sup>11</sup> AALTONEN 10K study  $p\bar{p} \rightarrow W^+W^-$  with  $W \rightarrow e/\mu\nu$ . The  $p_T$  of the leading (second) lepton is required to be  $> 20$  (10) GeV. The final number of events selected is 654 of which 320  $\pm$  47 are estimated to be background. The 95% C.L. interval is  $0.37 < \kappa_\gamma < 1.72$  for  $\Lambda = 1.5$  TeV and  $0.43 < \kappa_\gamma < 1.65$  for  $\Lambda = 2$  TeV.
- <sup>12</sup> AARON 09B study single- $W$  production in  $ep$  collisions at 0.3 TeV C.M. energy. They select 53  $W \rightarrow e/\mu$  events with a standard model expectation of  $54.1 \pm 7.4$  events. Fitting the transverse momentum spectrum of the hadronic recoil system they obtain a 95% C.L. limit of  $-3.7 < \kappa_\gamma < -1.5$  or  $0.3 < \kappa_\gamma < 1.5$ , where the ambiguity is due to the quadratic dependence of the cross section to the coupling parameter.
- <sup>13</sup> ABAZOV 09AD study the  $p\bar{p} \rightarrow \ell\nu 2\text{jet}$  process arising in  $WW$  and  $WZ$  production. They select 12,473 (14,392) events in the electron (muon) channel with an expected di-boson signal of 436 (527) events. The results on the anomalous couplings are derived from an analysis of the  $p_T$  spectrum of the 2-jet system and quoted at 68% C.L. and for a form factor of 2 TeV. This measurement is not used for obtaining the mean as it is for a specific form factor. The 95% confidence interval is  $0.56 < \kappa_\gamma < 1.55$ .
- <sup>14</sup> ABAZOV 09AJ study the  $p\bar{p} \rightarrow 2\ell 2\nu$  process arising in  $WW$  production. They select 100 events with an expected  $WW$  signal of 65 events. An analysis of the  $p_T$  spectrum of the two charged leptons leads to 95% C.L. limits of  $0.46 < \kappa_\gamma < 1.83$ , for a form factor  $\Lambda = 2$  TeV.
- <sup>15</sup> ABAZOV 08R use  $0.7 \text{ fb}^{-1}$   $p\bar{p}$  data at  $\sqrt{s} = 1.96$  TeV to select 263  $W\gamma + X$  events, of which 187 constitute signal, with the  $W$  decaying into an electron or a muon, which is required to be well separated from a photon with  $E_T > 9$  GeV. A likelihood fit to the photon  $E_T$  spectrum yields a 95% CL limit  $0.49 < \kappa_\gamma < 1.51$  with other couplings fixed to their Standard Model values.
- <sup>16</sup> ABDALLAH 08C determine this triple gauge coupling from the measurement of the spin density matrix elements in  $e^+e^- \rightarrow W^+W^- \rightarrow (q\bar{q})(\ell\nu)$ , where  $\ell = e$  or  $\mu$ . Values of all other couplings are fixed to their standard model values.
- <sup>17</sup> AALTONEN 07L set limits on anomalous TGCs using the  $p_T(W)$  distribution in  $WW$  and  $WZ$  production with the  $W$  decaying to an electron or muon and the  $Z$  to 2 jets. Setting other couplings to their standard model value, the 95% C.L. limits are  $0.54 < \kappa_\gamma < 1.39$  for a form factor scale  $\Lambda = 1.5$  TeV.
- <sup>18</sup> ABAZOV 06H study  $p\bar{p} \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+\nu_e e^-\bar{\nu}_e$ ,  $WW \rightarrow e^+\nu_e \mu^+\bar{\nu}_\mu$  or  $WW \rightarrow \mu^+\nu_\mu \mu^-\bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 1$  TeV is  $-0.05 < \kappa_\gamma < 2.29$ , fixing  $\lambda_\gamma = 0$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $0.68 < \kappa < 1.45$ .
- <sup>19</sup> ABAZOV 05J perform a likelihood fit to the photon  $E_T$  spectrum of  $W\gamma + X$  events, where the  $W$  decays to an electron or muon which is required to be well separated from the photon. For  $\Lambda = 2.0$  TeV the 95% CL limits are  $0.12 < \kappa_\gamma < 1.96$ . In the fit  $\lambda_\gamma$  is kept fixed to its Standard Model value.
- <sup>20</sup> ABREU 01I combine results from  $e^+e^-$  interactions at 189 GeV leading to  $W^+W^-$ ,  $W\nu_e$ , and  $\nu\bar{\nu}\gamma$  final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is  $0.87 < \kappa_\gamma < 1.68$ .
- <sup>21</sup> BREITWEG 00 search for  $W$  production in events with large hadronic  $p_T$ . For  $p_T > 20$  GeV, the upper limit on the cross section gives the 95%CL limit  $-3.7 < \kappa_\gamma < 2.5$  (for  $\lambda_\gamma = 0$ ).
- <sup>22</sup> ABBOTT 99I perform a simultaneous fit to the  $W\gamma$ ,  $WW \rightarrow$  dilepton,  $WW/WZ \rightarrow e\nu jj$ ,  $WW/WZ \rightarrow \mu\nu jj$ , and  $WZ \rightarrow$  trilepton data samples. For  $\Lambda = 2.0$  TeV, the 95%CL limits are  $0.75 < \kappa_\gamma < 1.39$ .

$\lambda_\gamma$

OUR FIT below is taken from [SCHAE 13A].

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.022 <math>\pm</math> 0.019 OUR FIT</b>				
0.002 $\pm$ 0.035	7872	<sup>1</sup> ABDALLAH	10 DLPH	$E_{\text{cm}}^e = 189\text{--}209$ GeV
-0.012 $\pm$ 0.027 $\pm$ 0.011	10689	<sup>2</sup> SCHAE	05A ALEP	$E_{\text{cm}}^e = 183\text{--}209$ GeV
-0.060 $\pm$ 0.034 -0.033	9800	<sup>3</sup> ABBIENDI	04D OPAL	$E_{\text{cm}}^e = 183\text{--}209$ GeV
-0.021 $\pm$ 0.035 $\pm$ 0.017	10575	<sup>4</sup> ACHARD	04D L3	$E_{\text{cm}}^e = 161\text{--}209$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

- <sup>5</sup> CHATRCHYAN 14AB CMS  $E_{\text{cm}}^{p\bar{p}} = 7$  TeV
- <sup>6</sup> AAD 13AN ATLS  $E_{\text{cm}}^{p\bar{p}} = 7$  TeV
- <sup>7</sup> ABAZOV 12AG D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>8</sup> ABAZOV 11AC D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>9</sup> CHATRCHYAN 11M CMS  $E_{\text{cm}}^{p\bar{p}} = 7$  TeV
- <sup>10</sup> AARON 09B H1  $E_{\text{cm}}^{ep} = 0.3$  TeV
- <sup>11</sup> ABAZOV 09AD D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>12</sup> ABAZOV 09AJ D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>13</sup> ABAZOV 08R D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>14</sup> ABDALLAH 08C DLPH Superseded by ABDALLAH 10
- <sup>15</sup> AALTONEN 07L CDF  $E_{\text{cm}}^{p\bar{p}} = 1.96$  GeV
- <sup>16</sup> ABAZOV 06H D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>17</sup> ABAZOV 05J D0  $E_{\text{cm}}^{p\bar{p}} = 1.96$  TeV
- <sup>18</sup> ABREU 01I DLPH  $E_{\text{cm}}^{ee} = 183\text{--}189$  GeV
- <sup>19</sup> BREITWEG 00 ZEUS  $e^+p \rightarrow e^+W^\pm X$ ,  $\sqrt{s} \approx 300$  GeV
- <sup>20</sup> ABBOTT 99I D0  $E_{\text{cm}}^{p\bar{p}} = 1.8$  TeV
- <sup>1</sup> ABDALLAH 10 use data on the final states  $e^+e^- \rightarrow jj\ell\nu, jjjj, jjX, \ell X$ , at center-of-mass energies between 189–209 GeV at LEP2, where  $j = \text{jet}$ ,  $\ell = \text{lepton}$ , and  $X$  represents missing momentum. The fit is carried out keeping all other parameters fixed at their SM values.
- <sup>2</sup> SCHAE 05A study single-photon, single- $W$ , and  $WW$ -pair production from 183 to 209 GeV. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- <sup>3</sup> ABBIENDI 04D combine results from  $W^+W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $-0.13 < \kappa_\gamma < 0.01$ .
- <sup>4</sup> ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- <sup>5</sup> CHATRCHYAN 14AB measure  $W\gamma$  production cross section for  $p_T^\gamma > 15$  GeV and  $R(\ell\gamma) > 0.7$ , which is the separation between the  $\gamma$  and the final state charged lepton ( $e$  or  $\mu$ ) in the azimuthal angle-pseudorapidity ( $\phi - \eta$ ) plane. After background subtraction the number of  $e\nu\gamma$  and  $\mu\nu\gamma$  events is determined to be  $3200 \pm 325$  and  $4970 \pm 543$  respectively, compatible with expectations from the SM. This leads to a 95% CL limit of  $-0.050 < \kappa_\gamma < 0.037$ , assuming all other parameters have SM values.
- <sup>6</sup> AAD 13AN study  $W\gamma$  production in  $pp$  collisions. In events with no additional jet, 4449 (6578)  $W$  decays to electron (muon) are selected, with an expected background of  $1662 \pm 262$  (2538  $\pm$  362) events. Analysing the photon  $p_T$  spectrum above 100 GeV yields a 95% C.L. limit of  $-0.065 < \kappa_\gamma < 0.061$ . Supersedes AAD 12Bx.
- <sup>7</sup> ABAZOV 12AG combine new results with already published results on  $W\gamma$ ,  $WW$  and  $WZ$  production in order to determine the couplings with increased precision, superseding ABAZOV 08R, ABAZOV 11AC, ABAZOV 09AJ, ABAZOV 09AD. The 68% C.L. result for a formfactor cutoff of  $\Lambda = 2$  TeV is  $\lambda_\gamma = 0.007^{+0.021}_{-0.022}$ .
- <sup>8</sup> ABAZOV 11AC study  $W\gamma$  production in  $p\bar{p}$  collisions at 1.96 TeV, with the  $W$  decay products containing an electron or a muon. They select 196 (363) events in the electron (muon) mode, with a SM expectation of 190 (372) events. A likelihood fit to the photon  $E_T$  spectrum above 15 GeV yields at 95% C.L. the result:  $-0.08 < \lambda_\gamma < 0.07$  for a formfactor  $\Lambda = 2$  TeV.
- <sup>9</sup> CHATRCHYAN 11M study  $W\gamma$  production in  $pp$  collisions at  $\sqrt{s} = 7$  TeV using 36 pb<sup>-1</sup>  $pp$  data with the  $W$  decaying to electron and muon. The total cross section is measured for photon transverse energy  $E_T^\gamma > 10$  GeV and spatial separation from charged leptons in the plane of pseudo rapidity and azimuthal angle  $\Delta R(\ell, \gamma) > 0.7$ . The number of candidate (background) events is 452 (228  $\pm$  21) for the electron channel and 520 (277  $\pm$  25) for the muon channel. Setting other couplings to their standard model value, they derive a 95% CL limit of  $-0.18 < \lambda_\gamma < 0.17$ .
- <sup>10</sup> AARON 09B study single- $W$  production in  $ep$  collisions at 0.3 TeV C.M. energy. They select 53  $W \rightarrow e/\mu$  events with a standard model expectation of  $54.1 \pm 7.4$  events. Fitting the transverse momentum spectrum of the hadronic recoil system they obtain a 95% C.L. limit of  $-2.5 < \lambda_\gamma < 2.5$ .
- <sup>11</sup> ABAZOV 09AD study the  $p\bar{p} \rightarrow \ell\nu 2\text{jet}$  process arising in  $WW$  and  $WZ$  production. They select 12,473 (14,392) events in the electron (muon) channel with an expected di-boson signal of 436 (527) events. The results on the anomalous couplings are derived from an analysis of the  $p_T$  spectrum of the 2-jet system and quoted at 68% C.L. and for a form factor of 2 TeV. This measurement is not used for obtaining the mean as it is for a specific form factor. The 95% confidence interval is  $-0.10 < \lambda_\gamma < 0.11$ .
- <sup>12</sup> ABAZOV 09AJ study the  $p\bar{p} \rightarrow 2\ell 2\nu$  process arising in  $WW$  production. They select 100 events with an expected  $WW$  signal of 65 events. An analysis of the  $p_T$  spectrum of the two charged leptons leads to 95% C.L. limits of  $-0.14 < \lambda_\gamma < 0.18$ , for a form factor  $\Lambda = 2$  TeV.
- <sup>13</sup> ABAZOV 08R use  $0.7 \text{ fb}^{-1}$   $p\bar{p}$  data at  $\sqrt{s} = 1.96$  TeV to select 263  $W\gamma + X$  events, of which 187 constitute signal, with the  $W$  decaying into an electron or a muon, which is required to be well separated from a photon with  $E_T > 9$  GeV. A likelihood fit to the photon  $E_T$  spectrum yields a 95% CL limit  $-0.12 < \lambda_\gamma < 0.13$  with other couplings fixed to their Standard Model values.
- <sup>14</sup> ABDALLAH 08C determine this triple gauge coupling from the measurement of the spin density matrix elements in  $e^+e^- \rightarrow W^+W^- \rightarrow (q\bar{q})(\ell\nu)$ , where  $\ell = e$  or  $\mu$ . Values of all other couplings are fixed to their standard model values.
- <sup>15</sup> AALTONEN 07L set limits on anomalous TGCs using the  $p_T(W)$  distribution in  $WW$  and  $WZ$  production with the  $W$  decaying to an electron or muon and the  $Z$  to 2 jets. Setting other couplings to their standard model value, the 95% C.L. limits are  $-0.18 < \lambda_\gamma < 0.17$  for a form factor scale  $\Lambda = 1.5$  TeV.

- <sup>16</sup>ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 1$  TeV is  $-0.97 < \lambda_\gamma < 1.04$ , fixing  $\kappa_\gamma=1$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $-0.29 < \lambda < 0.30$ .
- <sup>17</sup>ABAZOV 05J perform a likelihood fit to the photon  $E_T$  spectrum of  $W\gamma + X$  events, where the  $W$  decays to an electron or muon which is required to be well separated from the photon. For  $\Lambda = 2.0$  TeV the 95% CL limits are  $-0.20 < \lambda_\gamma < 0.20$ . In the fit  $\kappa_\gamma$  is kept fixed to its Standard Model value.
- <sup>18</sup>ABREU 01i combine results from  $e^+e^-$  interactions at 189 GeV leading to  $W^+W^-$ ,  $W\nu_e$ , and  $\nu\bar{\nu}\gamma$  final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is  $-0.11 < \lambda_\gamma < 0.23$ .
- <sup>19</sup>BREITWEG 00 search for  $W$  production in events with large hadronic  $p_T$ . For  $p_T > 20$  GeV, the upper limit on the cross section gives the 95%CL limit  $-3.2 < \lambda_\gamma < 3.2$  for  $\kappa_\gamma$  fixed to its Standard Model value.
- <sup>20</sup>ABBOTT 99i perform a simultaneous fit to the  $W\gamma$ ,  $WW \rightarrow$  dilepton,  $WW/WZ \rightarrow e\nu jj$ ,  $WW/WZ \rightarrow \mu\nu jj$ , and  $WZ \rightarrow$  trilepton data samples. For  $\Lambda = 2.0$  TeV, the 95%CL limits are  $-0.18 < \lambda_\gamma < 0.19$ .

 **$\kappa_Z$** 

This coupling is  $CP$ -conserving ( $C$ - and  $P$ - separately conserving).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.924 \pm 0.059$ $-0.056 \pm 0.024$	7171	<sup>1</sup> ACHARD 04D L3	$E_{cm}^{ee} = 189-209$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>2</sup> AAD 13AL ATLS	$E_{cm}^{pp} = 7$ TeV	
		<sup>3</sup> AAD 12CD ATLS	$E_{cm}^{pp} = 7$ TeV	
		<sup>4</sup> AALTONEN 12AC CDF	$E_{cm}^{pp} = 1.96$ TeV	
	34	<sup>5</sup> ABAZOV 11 D0	$E_{cm}^{pp} = 1.96$ TeV	
	17	<sup>6</sup> ABAZOV 06H D0	$E_{cm}^{pp} = 1.96$ TeV	
	2.3	<sup>7</sup> ABAZOV 05s D0	$E_{cm}^{pp} = 1.96$ TeV	

- <sup>1</sup>ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- <sup>2</sup>AAD 13AL study  $WW$  production in  $pp$  collisions and select 1325  $WW$  candidates in decay modes with electrons or muons with an expected background of  $369 \pm 61$  events. Assuming the LEP formulation and setting the form-factor  $\Lambda =$  infinity, a fit to the transverse momentum distribution of the leading charged lepton, leads to a 95% C.L. range of  $0.957 < \kappa_Z < 1.043$ . Supersedes AAD 12AC.
- <sup>3</sup>AAD 12CD study  $WZ$  production in  $pp$  collisions and select 317  $WZ$  candidates in three  $\ell\nu$  decay modes with an expected background of  $68.0 \pm 10.0$  events. The resulting 95% C.L. range is:  $0.63 < \kappa_Z < 1.57$ . Supersedes AAD 12V.
- <sup>4</sup>AALTONEN 12AC study  $WZ$  production in  $p\bar{p}$  collisions and select 63  $WZ$  candidates in three  $\ell\nu$  decay modes with an expected background of  $7.9 \pm 1.0$  events. Based on the cross section and shape of the  $Z$  transverse momentum spectrum, the following 95% C.L. range is reported:  $0.61 < \kappa_Z < 1.90$  for a form factor of  $\Lambda = 2$  TeV.
- <sup>5</sup>ABAZOV 11 study the  $p\bar{p} \rightarrow 3\ell\nu$  process arising in  $WZ$  production. They observe 34  $WZ$  candidates with an estimated background of 6 events. An analysis of the  $p_T$  spectrum of the  $Z$  boson leads to a 95% C.L. limit of  $0.600 < \kappa_Z < 1.675$ , for a form factor  $\Lambda = 2$  TeV.
- <sup>6</sup>ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 2$  TeV is  $0.55 < \kappa_Z < 1.55$ , fixing  $\lambda_Z=0$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $0.68 < \kappa < 1.45$ .
- <sup>7</sup>ABAZOV 05s study  $\bar{p}p \rightarrow WZ$  production with a subsequent trilepton decay to  $\ell\nu\ell'\bar{\ell}'$  ( $\ell$  and  $\ell' = e$  or  $\mu$ ). Three events (estimated background  $0.71 \pm 0.08$  events) with  $WZ$  decay characteristics are observed from which they derive limits on the anomalous  $WWZ$  couplings. The 95% CL limit for a form factor scale  $\Lambda = 1$  TeV is  $-1.0 < \kappa_Z < 3.4$ , fixing  $\lambda_Z$  and  $g_1^Z$  to their Standard Model values.

 **$\lambda_Z$** 

This coupling is  $CP$ -conserving ( $C$ - and  $P$ - separately conserving).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.088 \pm 0.060$ $-0.057 \pm 0.023$	7171	<sup>1</sup> ACHARD 04D L3	$E_{cm}^{ee} = 189-209$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>2</sup> AAD 14Y ATLS	$E_{cm}^{pp} = 8$ TeV	
		<sup>3</sup> AAD 13AL ATLS	$E_{cm}^{pp} = 7$ TeV	
		<sup>4</sup> CHATRCHYAN 13BF CMS	$E_{cm}^{pp} = 7$ TeV	
		<sup>5</sup> AAD 12CD ATLS	$E_{cm}^{pp} = 7$ TeV	
		<sup>6</sup> AALTONEN 12AC CDF	$E_{cm}^{pp} = 1.96$ TeV	
	34	<sup>7</sup> ABAZOV 11 D0	$E_{cm}^{pp} = 1.96$ TeV	
	334	<sup>8</sup> AALTONEN 10K CDF	$E_{cm}^{pp} = 1.96$ TeV	
	13	<sup>9</sup> ABAZOV 07Z D0	$E_{cm}^{pp} = 1.96$ TeV	
	17	<sup>10</sup> ABAZOV 06H D0	$E_{cm}^{pp} = 1.96$ TeV	
	2.3	<sup>11</sup> ABAZOV 05s D0	$E_{cm}^{pp} = 1.96$ TeV	

- <sup>1</sup>ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- <sup>2</sup>AAD 14Y determine the electroweak  $Z$ -dijet cross section in 8 TeV  $pp$  collisions.  $Z \rightarrow ee$  and  $Z \rightarrow \mu\mu$  decays are selected with the di-lepton  $p_T > 20$  GeV and mass in the 81-101 GeV range. Minimum two jets are required with  $p_T > 55$  and 45 GeV and no additional jets with  $p_T > 25$  GeV in the rapidity interval between them. The normalized  $p_T$  balance between the  $Z$  and the two jets is required to be  $< 0.15$ . This leads to a selection of 900 events with dijet mass  $> 1$  TeV. The number of signal and background events expected is 261 and 592 respectively. A Poisson likelihood method is used on an event by event basis to obtain the 95% CL limit  $-0.15 < \lambda_Z < 0.13$  for a form factor value  $\Lambda = \infty$ .
- <sup>3</sup>AAD 13AL study  $WW$  production in  $pp$  collisions and select 1325  $WW$  candidates in decay modes with electrons or muons with an expected background of  $369 \pm 61$  events. Assuming the LEP formulation and setting the form-factor  $\Lambda =$  infinity, a fit to the transverse momentum distribution of the leading charged lepton, leads to a 95% C.L. range of  $-0.062 < \lambda_Z < 0.059$ . Supersedes AAD 12AC.
- <sup>4</sup>CHATRCHYAN 13BF determine the  $W^+W^-$  production cross section using unlike sign di-lepton ( $e$  or  $\mu$ ) events with high  $p_T$ . The leptons have  $p_T > 20$  GeV/c and are isolated. 1134 candidate events are observed with an expected SM background of  $247 \pm 34$ . The  $p_T$  distribution of the leading lepton is fitted to obtain 95% C.L. limits of  $-0.048 \leq \lambda_Z \leq 0.048$ .
- <sup>5</sup>AAD 12CD study  $WZ$  production in  $pp$  collisions and select 317  $WZ$  candidates in three  $\ell\nu$  decay modes with an expected background of  $68.0 \pm 10.0$  events. The resulting 95% C.L. range is:  $-0.046 < \lambda_Z < 0.047$ . Supersedes AAD 12V.
- <sup>6</sup>AALTONEN 12AC study  $WZ$  production in  $p\bar{p}$  collisions and select 63  $WZ$  candidates in three  $\ell\nu$  decay modes with an expected background of  $7.9 \pm 1.0$  events. Based on the cross section and shape of the  $Z$  transverse momentum spectrum, the following 95% C.L. range is reported:  $-0.08 < \lambda_Z < 0.10$  for a form factor of  $\Lambda = 2$  TeV.
- <sup>7</sup>ABAZOV 11 study the  $p\bar{p} \rightarrow 3\ell\nu$  process arising in  $WZ$  production. They observe 34  $WZ$  candidates with an estimated background of 6 events. An analysis of the  $p_T$  spectrum of the  $Z$  boson leads to a 95% C.L. limit of  $-0.077 < \lambda_Z < 0.093$ , for a form factor  $\Lambda = 2$  TeV.
- <sup>8</sup>AALTONEN 10K study  $p\bar{p} \rightarrow W^+W^-$  with  $W \rightarrow e/\mu\nu$ . The  $p_T$  of the leading (second) lepton is required to be  $> 20$  (10) GeV. The final number of events selected is 654 of which  $320 \pm 47$  are estimated to be background. The 95% C.L. interval is  $-0.16 < \lambda_Z < 0.16$  for  $\Lambda = 1.5$  TeV and  $-0.14 < \lambda_Z < 0.15$  for  $\Lambda = 2$  TeV.
- <sup>9</sup>ABAZOV 07Z set limits on anomalous TGCs using the measured cross section and  $p_T(Z)$  distribution in  $WZ$  production with both the  $W$  and the  $Z$  decaying leptonically into electrons and muons. Setting the other couplings to their standard model values, the 95% C.L. limit for a form factor scale  $\Lambda = 2$  TeV is  $-0.17 < \lambda_Z < 0.21$ .
- <sup>10</sup>ABAZOV 06H study  $\bar{p}p \rightarrow WW$  production with a subsequent decay  $WW \rightarrow e^+ \nu_e e^- \bar{\nu}_e$ ,  $WW \rightarrow e^\pm \nu_e \mu^\mp \nu_\mu$  or  $WW \rightarrow \mu^+ \nu_\mu \mu^- \bar{\nu}_\mu$ . The 95% C.L. limit for a form factor scale  $\Lambda = 2$  TeV is  $-0.39 < \lambda_Z < 0.39$ , fixing  $\kappa_Z=1$ . With the assumption that the  $WW\gamma$  and  $WWZ$  couplings are equal the 95% C.L. one-dimensional limit ( $\Lambda = 2$  TeV) is  $-0.29 < \lambda < 0.30$ .
- <sup>11</sup>ABAZOV 05s study  $\bar{p}p \rightarrow WZ$  production with a subsequent trilepton decay to  $\ell\nu\ell'\bar{\ell}'$  ( $\ell$  and  $\ell' = e$  or  $\mu$ ). Three events (estimated background  $0.71 \pm 0.08$  events) with  $WZ$  decay characteristics are observed from which they derive limits on the anomalous  $WWZ$  couplings. The 95% CL limit for a form factor scale  $\Lambda = 1.5$  TeV is  $-0.48 < \lambda_Z < 0.48$ , fixing  $g_1^Z$  and  $\kappa_Z$  to their Standard Model values.

 **$g_5^Z$** 

This coupling is  $CP$ -conserving but  $C$ - and  $P$ -violating.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.07 \pm 0.09$ OUR AVERAGE		Error includes scale factor of 1.1.		
$-0.04 \pm 0.13$ $-0.12$	9800	<sup>1</sup> ABBIENDI 04D OPAL	$E_{cm}^{ee} = 183-209$ GeV	
$0.00 \pm 0.13 \pm 0.05$	7171	<sup>2</sup> ACHARD 04D L3	$E_{cm}^{ee} = 189-209$ GeV	
$-0.44 \pm 0.23 \pm 0.12$	1154	<sup>3</sup> ACCIARRI 99Q L3	$E_{cm}^{ee} = 161+172+183$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.31 \pm 0.23$		<sup>4</sup> EBOLI 00 THEO LEP1, SLAC+ Tevatron		

- <sup>1</sup>ABBIENDI 04D combine results from  $W^+W^-$  in all decay channels. Only  $CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is  $-0.28 < g_5^Z < +0.21$ .
- <sup>2</sup>ACHARD 04D study  $WW$ -pair production, single- $W$  production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the  $WW$ -pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- <sup>3</sup>ACCIARRI 99Q study  $W$ -pair, single- $W$ , and single photon events.
- <sup>4</sup>EBOLI 00 extract this indirect value of the coupling studying the non-universal one-loop contributions to the experimental value of the  $Z \rightarrow b\bar{b}$  width ( $\Lambda=1$  TeV is assumed).

 **$g_4^Z$** 

This coupling is  $CP$ -violating ( $C$ -violating and  $P$ -conserving).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.30 \pm 0.17$ OUR AVERAGE				
$-0.39 \pm 0.19$ $-0.20$	1880	<sup>1</sup> ABDALLAH 08c DLPH	$E_{cm}^{ee} = 189-209$ GeV	
$-0.02 \pm 0.32$ $-0.33$	1065	<sup>2</sup> ABBIENDI 01H OPAL	$E_{cm}^{ee} = 189$ GeV	

- <sup>1</sup>ABDALLAH 08c determine this triple gauge coupling from the measurement of the spin density matrix elements in  $e^+e^- \rightarrow W^+W^- \rightarrow (q\bar{q})(\ell\nu)$ , where  $\ell = e$  or  $\mu$ . Values of all other couplings are fixed to their standard model values.
- <sup>2</sup>ABBIENDI 01H study  $W$ -pair events, with one leptonically and one hadronically decaying  $W$ . The coupling is extracted using information from the  $W$  production angle together with decay angles from the leptonically decaying  $W$ .

# Gauge & Higgs Boson Particle Listings

## W

### $\tilde{\kappa}_Z$

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.12^{+0.06}_{-0.04}$ OUR AVERAGE				
$-0.09^{+0.08}_{-0.05}$	1880	<sup>1</sup> ABDALLAH	08c DLPH	$E_{\text{cm}}^{\text{ee}} = 189\text{--}209$ GeV
$-0.20^{+0.10}_{-0.07}$	1065	<sup>2</sup> ABBIENDI	01H OPAL	$E_{\text{cm}}^{\text{ee}} = 189$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>3</sup> BLINOV	11 LEP	$E_{\text{cm}}^{\text{ee}} = 183\text{--}207$ GeV

<sup>1</sup> ABDALLAH 08c determine this triple gauge coupling from the measurement of the spin density matrix elements in  $e^+e^- \rightarrow W^+W^- \rightarrow (qq)(\ell\nu)$ , where  $\ell = e$  or  $\mu$ . Values of all other couplings are fixed to their standard model values.

<sup>2</sup> ABBIENDI 01H study  $W$ -pair events, with one leptonically and one hadronically decaying  $W$ . The coupling is extracted using information from the  $W$  production angle together with decay angles from the leptonically decaying  $W$ .

<sup>3</sup> BLINOV 11 use the LEP-average  $e^+e^- \rightarrow W^+W^-$  cross section data for  $\sqrt{s} = 183\text{--}207$  GeV to determine an upper limit on the TGC  $\tilde{\kappa}_Z$ . The average values of the cross sections as well as their correlation matrix, and standard model expectations of the cross sections are taken from the LEPEWWG note hep-ex/0612034. At 95% confidence level  $|\tilde{\kappa}_Z| < 0.13$ .

### $\tilde{\lambda}_Z$

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.09 \pm 0.07$ OUR AVERAGE				
$-0.08 \pm 0.07$	1880	<sup>1</sup> ABDALLAH	08c DLPH	$E_{\text{cm}}^{\text{ee}} = 189\text{--}209$ GeV
$-0.18^{+0.24}_{-0.16}$	1065	<sup>2</sup> ABBIENDI	01H OPAL	$E_{\text{cm}}^{\text{ee}} = 189$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>3</sup> BLINOV	11 LEP	$E_{\text{cm}}^{\text{ee}} = 183\text{--}207$ GeV

<sup>1</sup> ABDALLAH 08c determine this triple gauge coupling from the measurement of the spin density matrix elements in  $e^+e^- \rightarrow W^+W^- \rightarrow (qq)(\ell\nu)$ , where  $\ell = e$  or  $\mu$ . Values of all other couplings are fixed to their standard model values.

<sup>2</sup> ABBIENDI 01H study  $W$ -pair events, with one leptonically and one hadronically decaying  $W$ . The coupling is extracted using information from the  $W$  production angle together with decay angles from the leptonically decaying  $W$ .

<sup>3</sup> BLINOV 11 use the LEP-average  $e^+e^- \rightarrow W^+W^-$  cross section data for  $\sqrt{s} = 183\text{--}207$  GeV to determine an upper limit on the TGC  $\tilde{\lambda}_Z$ . The average values of the cross sections as well as their correlation matrix, and standard model expectations of the cross sections are taken from the LEPEWWG note hep-ex/0612034. At 95% confidence level  $|\tilde{\lambda}_Z| < 0.31$ .

## W ANOMALOUS MAGNETIC MOMENT

The full magnetic moment is given by  $\mu_W = e(1 + \kappa + \lambda)/2m_W$ . In the Standard Model, at tree level,  $\kappa = 1$  and  $\lambda = 0$ . Some papers have defined  $\Delta\kappa = 1 - \kappa$  and assume that  $\lambda = 0$ . Note that the electric quadrupole moment is given by  $-e(\kappa - \lambda)/m_W^2$ . A description of the parameterization of these moments and additional references can be found in HAGIWARA 87 and BAUR 88. The parameter  $\Lambda$  appearing in the theoretical limits below is a regularization cutoff which roughly corresponds to the energy scale where the structure of the  $W$  boson becomes manifest.

VALUE ( $e/2m_W$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$2.22^{+0.20}_{-0.19}$	2298	<sup>1</sup> ABREU	01i DLPH	$E_{\text{cm}}^{\text{ee}} = 183\text{--}189$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>2</sup> ABE	95g CDF	
		<sup>3</sup> ALITTI	92c UA2	
		<sup>4</sup> SAMUEL	92 THEO	
		<sup>5</sup> SAMUEL	91 THEO	
		<sup>6</sup> GRIFOLS	88 THEO	
		<sup>7</sup> GROTH	87 THEO	
		<sup>8</sup> VANDERBIJ	87 THEO	
		<sup>9</sup> GRAU	85 THEO	
		<sup>10</sup> SUZUKI	85 THEO	
		<sup>11</sup> HERZOG	84 THEO	

<sup>1</sup> ABREU 01i combine results from  $e^+e^-$  interactions at 189 GeV leading to  $W^+W^-$ ,  $W\nu_e$ , and  $\nu\bar{\nu}\gamma$  final states with results from ABREU 99L at 183 GeV to determine  $\Delta g_Z^2$ ,  $\Delta\kappa_\gamma$ , and  $\lambda_\gamma$ .  $\Delta\kappa_\gamma$  and  $\lambda_\gamma$  are simultaneously floated in the fit to determine  $\mu_W$ .

<sup>2</sup> ABE 95g report  $-1.3 < \kappa < 3.2$  for  $\lambda = 0$  and  $-0.7 < \lambda < 0.7$  for  $\kappa = 1$  in  $p\bar{p} \rightarrow e\nu_e\gamma X$  and  $\mu\nu_\mu\gamma X$  at  $\sqrt{s} = 1.8$  TeV.

<sup>3</sup> ALITTI 92c measure  $\kappa = 1 \pm 2.6$  and  $\lambda = 0 \pm 1.7$  in  $p\bar{p} \rightarrow e\nu\gamma + X$  at  $\sqrt{s} = 630$  GeV. At 95%CL they report  $-3.5 < \kappa < 5.9$  and  $-3.6 < \lambda < 3.5$ .

<sup>4</sup> SAMUEL 92 use preliminary CDF and UA2 data and find  $-2.4 < \kappa < 3.7$  at 96%CL and  $-3.1 < \kappa < 4.2$  at 95%CL respectively. They use data for  $W\gamma$  production and radiative  $W$  decay.

<sup>5</sup> SAMUEL 91 use preliminary CDF data for  $p\bar{p} \rightarrow W\gamma X$  to obtain  $-11.3 \leq \Delta\kappa \leq 10.9$ . Note that their  $\kappa = 1 - \Delta\kappa$ .

<sup>6</sup> GRIFOLS 88 uses deviation from  $\rho$  parameter to set limit  $\Delta\kappa \lesssim 65 (M_W^2/\Lambda^2)$ .

<sup>7</sup> GROTH 87 finds the limit  $-37 < \Delta\kappa < 73.5$  (90% CL) from the experimental limits on  $e^+e^- \rightarrow \nu\bar{\nu}\gamma$  assuming three neutrino generations and  $-19.5 < \Delta\kappa < 56$  for four generations. Note their  $\Delta\kappa$  has the opposite sign as our definition.

<sup>8</sup> VANDERBIJ 87 uses existing limits to the photon structure to obtain  $|\Delta\kappa| < 33 (m_W/\Lambda)$ . In addition VANDERBIJ 87 discusses problems with using the  $\rho$  parameter of the Standard Model to determine  $\Delta\kappa$ .

<sup>9</sup> GRAU 85 uses the muon anomaly to derive a coupled limit on the anomalous magnetic dipole and electric quadrupole ( $\lambda$ ) moments  $1.05 > \Delta\kappa \ln(\Lambda/m_W) + \lambda/2 > -2.77$ . In the Standard Model  $\lambda = 0$ .

<sup>10</sup> SUZUKI 85 uses partial-wave unitarity at high energies to obtain  $|\Delta\kappa| \lesssim 190 (m_W/\Lambda)^2$ . From the anomalous magnetic moment of the muon, SUZUKI 85 obtains  $|\Delta\kappa| \lesssim 2.2/\ln(\Lambda/m_W)$ . Finally SUZUKI 85 uses deviations from the  $\rho$  parameter and obtains a very qualitative, order-of-magnitude limit  $|\Delta\kappa| \lesssim 150 (m_W/\Lambda)^4$  if  $|\Delta\kappa| \ll 1$ .

<sup>11</sup> HERZOG 84 consider the contribution of  $W$ -boson to muon magnetic moment including anomalous coupling of  $WW\gamma$ . Obtain a limit  $-1 < \Delta\kappa < 3$  for  $\Lambda \gtrsim 1$  TeV.

## ANOMALOUS W/Z QUARTIC COUPLINGS

### ANOMALOUS W/Z QUARTIC COUPLINGS (QGCS)

Revised November 2015 by M.W. Gr newald (U. College Dublin) and A. Gurtu (Formerly Tata Inst.).

Quartic couplings,  $WWZZ$ ,  $WWZ\gamma$ ,  $WW\gamma\gamma$ , and  $ZZ\gamma\gamma$ , were studied at LEP and Tevatron at energies at which the Standard Model predicts negligible contributions to multiboson production. Thus, to parametrize limits on these couplings, an effective theory approach is adopted which supplements the Standard Model Lagrangian with higher dimensional operators which include quartic couplings. The LEP collaborations chose the lowers dimensional representation of operators (dimension 6) which presumes the  $SU(2) \times U(1)$  gauge symmetry is broken by means other than the conventional Higgs scalar doublet [1–3]. In this representation possible quartic couplings,  $a_0, a_c, a_n$ , are expressed in terms of the following dimension-6 operators [1,2];

$$\begin{aligned} L_6^0 &= -\frac{e^2}{16\Lambda^2} a_0 F^{\mu\nu} F_{\mu\nu} \vec{W}^\alpha \cdot \vec{W}_\alpha \\ L_6^c &= -\frac{e^2}{16\Lambda^2} a_c F^{\mu\alpha} F_{\mu\beta} \vec{W}^\beta \cdot \vec{W}_\alpha \\ L_6^n &= -i\frac{e^2}{16\Lambda^2} a_n \epsilon_{ijk} W_{\mu\alpha}^{(i)} W_{\nu}^{(j)} W^{(k)\alpha} F^{\mu\nu} \\ \tilde{L}_6^0 &= -\frac{e^2}{16\Lambda^2} \tilde{a}_0 F^{\mu\nu} \tilde{F}_{\mu\nu} \vec{W}^\alpha \cdot \vec{W}_\alpha \\ \tilde{L}_6^n &= -i\frac{e^2}{16\Lambda^2} \tilde{a}_n \epsilon_{ijk} \tilde{W}_{\mu\alpha}^{(i)} W_{\nu}^{(j)} W^{(k)\alpha} \tilde{F}^{\mu\nu} \end{aligned}$$

where  $F, W$  are photon and  $W$  fields,  $L_6^0$  and  $L_6^c$  conserve  $C$ ,  $P$  separately ( $\tilde{L}_6^0$  conserves only  $C$ ) and generate anomalous  $W^+W^-\gamma\gamma$  and  $ZZ\gamma\gamma$  couplings,  $L_6^n$  violates  $CP$  ( $\tilde{L}_6^n$  violates both  $C$  and  $P$ ) and generates an anomalous  $W^+W^-Z\gamma$  coupling, and  $\Lambda$  is an energy scale for new physics. For the  $ZZ\gamma\gamma$  coupling the  $CP$ -violating term represented by  $L_6^n$  does not contribute. These couplings are assumed to be real and to vanish at tree level in the Standard Model.

Within the same framework as above, a more recent description of the quartic couplings [3] treats the anomalous parts of the  $WW\gamma\gamma$  and  $ZZ\gamma\gamma$  couplings separately, leading to two sets parametrized as  $a_0^V/\Lambda^2$  and  $a_c^V/\Lambda^2$ , where  $V = W$  or  $Z$ .

With the discovery of a Higgs at the LHC in 2012, it is then useful to go to the next higher dimensional representation (dimension 8 operators) in which the gauge symmetry is broken by the conventional Higgs scalar doublet [3,4]. There are 14 operators which can contribute to the anomalous quartic coupling signal. Some of the operators have analogues in the dimension 6 scheme. The CMS collaboration, [5], have used

this parametrization, in which the connections between the two schemes are also summarized:

$$\begin{aligned}\mathcal{L}_{AQGC} = & -\frac{e^2 a_0^W}{8 \Lambda^2} F_{\mu\nu} F^{\mu\nu} W^{+a} W_a^- \\ & -\frac{e^2 a_c^W}{16 \Lambda^2} F_{\mu\nu} F^{\mu a} (W^{+\nu} W_a^- + W^{-\nu} W_a^+) \\ & -e^2 g^2 \frac{\kappa_0^W}{\Lambda^2} F_{\mu\nu} Z^{\mu\nu} W^{+a} W_a^- \\ & -\frac{e^2 g^2 \kappa_c^W}{2 \Lambda^2} F_{\mu\nu} Z^{\mu a} (W^{+\nu} W_a^- + W^{-\nu} W_a^+) \\ & + \frac{f_{T,0}}{\Lambda^4} \text{Tr}[\widehat{W}_{\mu\nu} \widehat{W}^{\mu\nu}] \times \text{Tr}[\widehat{W}_{\alpha\beta} \widehat{W}^{\alpha\beta}]\end{aligned}$$

The energy scale of possible new physics is  $\Lambda$ , and  $g = e/\sin(\theta_W)$ ,  $e$  being the unit electric charge and  $\theta_W$  the Weinberg angle. The field tensors are described in [3,4].

The two dimension 6 operators  $a_0^W/\Lambda^2$  and  $a_c^W/\Lambda^2$  are associated with the  $WW\gamma\gamma$  vertex. Among dimension 8 operators,  $\kappa_0^W/\Lambda^2$  and  $\kappa_c^W/\Lambda^2$  are associated with the  $WWZ\gamma$  vertex, whereas the parameter  $f_{T,0}/\Lambda^4$  contributes to both vertices. There is a relationship between these two dimension 6 parameters and the dimension 8 parameters  $f_{M,i}/\Lambda^4$  as follows [3]:

$$\begin{aligned}\frac{a_0^W}{\Lambda^2} &= -\frac{4M_W^2}{g^2} \frac{f_{M,0}}{\Lambda^4} - \frac{8M_W^2}{g'^2} \frac{f_{M,2}}{\Lambda^4} \\ \frac{a_c^W}{\Lambda^2} &= -\frac{4M_W^2}{g^2} \frac{f_{M,1}}{\Lambda^4} - \frac{8M_W^2}{g'^2} \frac{f_{M,3}}{\Lambda^4}\end{aligned}$$

where  $g' = e/\cos(\theta_W)$  and  $M_W$  is the invariant mass of the  $W$  boson. This relation provides a translation between limits on dimension 6 operators  $a_{0,c}^W$  and  $f_{M,j}/\Lambda^4$ . It is further required [4] that  $f_{M,0} = 2f_{M,2}$  and  $f_{M,1} = 2f_{M,3}$  which suppresses contributions to the  $WWZ\gamma$  vertex. The complete set of Lagrangian contributions as presented in [4] corresponds to 19 anomalous couplings in total –  $f_{S,i}$ ,  $i = 1, 2$ ,  $f_{M,i}$ ,  $i = 0, \dots, 8$  and  $f_{T,i}$ ,  $i = 0, \dots, 9$  – each scaled by  $1/\Lambda^4$ .

The ATLAS collaboration [6], on the other hand, follows a K-matrix driven approach of Ref. 7 in which the anomalous couplings can be expressed in terms of two parameters  $\alpha_4$  and  $\alpha_5$ , which account for all BSM effects.

It is the early stages in the determination of quartic couplings by the LHC experiments. It is hoped that the two collaborations, ATLAS and CMS, will agree to use at least one common set of parameters to express these limits to enable the reader to make a comparison and allow for a possible LHC combination.

## References

1. G. Belanger and F. Boudjema, Phys. Lett. **B288**, 201 (1992).
2. J.W. Stirling and A. Werthenbach, Eur. Phys. J. **C14**, 103 (2000);  
J.W. Stirling and A. Werthenbach, Phys. Lett. **B466**, 369 (1999);  
A. Denner *et al.*, Eur. Phys. J. **C20**, 201 (2001);  
G. Montagna *et al.*, Phys. Lett. **B515**, 197 (2001).
3. G. Belanger *et al.*, Eur. Phys. J. **C13**, 283 (2000).

4. O.J.P. Éboli, M.C. Gonzalez-Garcia, and S.M. Lietti, Phys. Rev. **D69**, 095005 (2004);  
O.J.P. Éboli, M.C. Gonzalez-Garcia, and J.K. Mizukoshi, Phys. Rev. **D77**, 073005 (2006).
5. S. Chatrchyan *et al.*, Phys. Rev. **D90**, 032008 (2014);  
S. Chatrchyan *et al.*, Phys. Rev. Lett. **114**, 051801 (2015).
6. G. Aad *et al.*, Phys. Rev. Lett. **113**, 141803 (2014).
7. A. Albateanu, W. Killian, and J. Reuter, JHEP **0811**, 010 (2008).

$a_0/\Lambda^2, a_c/\Lambda^2, a_n/\Lambda^2, \kappa_0^W/\Lambda^2, \kappa_c^W/\Lambda^2, f_{T,0}/\Lambda^4, f_{M,i}/\Lambda^4, \alpha_4, \alpha_5, f_{S,i}/\Lambda^4, f_{M,i}/\Lambda^4, f_{T,i}/\Lambda^4$

Anomalous  $W$  quartic couplings are measured by the experiments at LEP, the Tevatron, and the LHC. Some of the recent results from the Tevatron and LHC experiments individually surpass the combined LEP-2 results in precision (see below). As discussed in the review on the “Anomalous  $W/Z$  quartic couplings (QGCS),” the measurements are typically done using different operator expansions which then do not allow the results to be compared and averaged. At least one common framework should be agreed upon for the use in the future publications by the experiments.

VALUE DOCUMENT ID TECN  
• • • We do not use the following data for averages, fits, limits, etc. • • •

1	AAD	15N	ATLS
2	KHACHATRYAN	15D	CMS
3	AAD	14AM	ATLS
4	CHATRCHYAN	14Q	CMS
5	ABAZOV	13D	D0
6	CHATRCHYAN	13AA	CMS
7	ABBIENDI	04B	OPAL
8	ABBIENDI	04L	OPAL
9	HEISTER	04A	ALEP
10	ABDALLAH	03I	DLPH
11	ACHARD	02F	L3

<sup>1</sup>AAD 15N study  $W\gamma\gamma$  events in 8 TeV  $pp$  interactions, where the  $W$  decays into an electron or a muon. The events are characterized by an isolated lepton, a missing transverse energy due to the decay neutrino, and two isolated photons, with the  $p_T$  of the lepton and the photons being  $> 20$  GeV. The number of candidate events observed in the electron channel for  $N(\text{jet}) \geq 0$  and  $N(\text{jet}) = 0$  is 47 and 15, the corresponding numbers for the muon channel being 110 and 53. The backgrounds expected are  $30.2 \pm 7.4$ ,  $8.7 \pm 3.0$ ,  $52.1 \pm 12.2$ , and  $24.4 \pm 8.3$  respectively. The 95% C.L. limits on the values of the parameters  $f_{T,0}/\Lambda^4$ ,  $f_{M,2}/\Lambda^4$  and  $f_{M,3}/\Lambda^4$  are  $-0.9$ – $0.9 \times 10^2$ ,  $-0.8$ – $0.8 \times 10^4$ , and  $-1.5$ – $1.4 \times 10^4$  respectively, without application of a form factor  $\Lambda_{FF}$ .

<sup>2</sup>KHACHATRYAN 15D study vector-boson-scattering tagged by two jets, requiring two same-sign charged leptons arising from  $W^\pm W^\pm$  production and decay. The two jets must have a transverse momentum larger than 30 GeV, while the leptons, electrons or muons, must have a transverse momentum  $> 20$  GeV. The dijet mass is required to be  $> 500$  GeV, the dilepton mass  $> 50$  GeV, with additional requirement of differing from the  $Z$  mass by  $> 15$  GeV. In the two categories  $W^+ W^+$  and  $W^- W^-$ , 10 and 2 data events are observed in a data sample corresponding to an integrated luminosity of  $19.4 \text{ fb}^{-1}$ , with an expected background of  $3.1 \pm 0.6$  and  $2.6 \pm 0.5$  events. Analysing the distribution of the dilepton invariant mass, the following limits at 95% C.L. are obtained, in units of  $\text{TeV}^{-4}$ :  $-38 < f_{S,0}/\Lambda^4 < 40$ ,  $-118 < f_{S,1}/\Lambda^4 < 120$ ,  $-33 < f_{M,0}/\Lambda^4 < 32$ ,  $-44 < f_{M,1}/\Lambda^4 < 47$ ,  $-65 < f_{M,6}/\Lambda^4 < 63$ ,  $-70 < f_{M,7}/\Lambda^4 < 66$ ,  $-4.2 < f_{T,0}/\Lambda^4 < 4.6$ ,  $-1.9 < f_{T,1}/\Lambda^4 < 2.2$ ,  $-6.2 < f_{T,2}/\Lambda^4 < 6.4$ .

<sup>3</sup>AAD 14AM analyze electroweak production of  $WW$  jet same-charge diboson plus two jets production, with the  $W$  bosons decaying to electron or muon, to study the quartic  $WWWW$  coupling. In a kinematic region enhancing the electroweak production over the strong production, 34 events are observed in the data while  $29.8 \pm 2.4$  events are expected with a background of  $15.9 \pm 1.9$  events. Assuming the other QGC coupling to have the SM value of zero, the observed event yield is used to determine 95% CL limits on the quartic gauge couplings:  $-0.14 < \alpha_4 < 0.16$  and  $-0.23 < \alpha_5 < 0.24$ .

<sup>4</sup>CHATRCHYAN 14Q study  $WV\gamma$  production in 8 TeV  $pp$  collisions, in the single lepton final state, with  $W \rightarrow \ell\nu$ ,  $Z \rightarrow \text{dijet}$  or  $W \rightarrow \ell\nu$ ,  $W \rightarrow \text{dijet}$ , the dijet mass resolution precluding differentiation between the  $W$  and  $Z$ .  $p_T$  and pseudo-rapidity cuts are put on the lepton, the photon and the two jets to minimize backgrounds. The dijet mass is required to be between 7–100 GeV and  $|\Delta\eta_{jj}| < 1.4$ . The selected number of muon (electron) events are 183 (139), with SM expectation being  $194.2 \pm 11.5$  ( $147.9 \pm 10.7$ ) including signal and background. The photon  $E_T$  distribution is used to set limits on the anomalous quartic couplings. The following 95% CL limits are deduced (all in units of  $\text{TeV}^{-2}$  or  $\text{TeV}^{-4}$ ):  $-21 < a_0^W/\Lambda^2 < 20$ ,  $-34 < a_c^W/\Lambda^2 < 32$ ,  $-12 < \kappa_0^W/\Lambda^2 < 10$  and  $-18 < \kappa_c^W/\Lambda^2 < 17$ ; and  $-25 < f_{T,0}/\Lambda^4 < 24 \text{ TeV}^{-4}$ .

<sup>5</sup>ABAZOV 13D searches for anomalous  $WW\gamma\gamma$  quartic gauge couplings in the two-photon-mediated process  $pp \rightarrow ppWW$ , assuming the  $WW\gamma$  triple gauge boson couplings to be at their Standard Model values. 946 events containing an  $e^+e^-$  pair with missing energy are selected in a total luminosity of  $9.7 \text{ fb}^{-1}$ , with an expectation of  $983 \pm 108$  events from Standard-Model processes. The following 1-parameter limits at 95% CL are obtained:  $|a_0^W/\Lambda^2| < 4.3 \times 10^{-4} \text{ GeV}^{-2}$  ( $a_c^W = 0$ ),  $|a_c^W/\Lambda^2| < 1.5 \times 10^{-3} \text{ GeV}^{-2}$  ( $a_0^W = 0$ ).

<sup>6</sup>CHATRCHYAN 13AA searches for anomalous  $WW\gamma\gamma$  quartic gauge couplings in the two-photon-mediated process  $pp \rightarrow ppWW$ , assuming the  $WW\gamma$  triple gauge boson couplings to be at their Standard Model values. 2 events containing an  $e^\pm\mu^\mp$  pair with  $p_T(e, \mu) > 30$  GeV are selected in a total luminosity of  $5.05 \text{ fb}^{-1}$ , with an expected  $ppWW$  signal of  $2.2 \pm 0.4$  events and an expected background of  $0.84 \pm 0.15$  events. The following 1-parameter limits at 95% CL are obtained from the  $p_T(e, \mu)$  spectrum:

# Gauge & Higgs Boson Particle Listings

## W, Z

$|a_0^W/\Lambda^2| < 4.0 \times 10^{-6} \text{ GeV}^{-2}$  ( $a_c^W = 0$ ),  $|a_c^W/\Lambda^2| < 1.5 \times 10^{-5} \text{ GeV}^{-2}$  ( $a_0^W = 0$ ).

<sup>7</sup> ABBIENDI 04B select  $187 e^+e^- \rightarrow W^+W^-\gamma$  events in the C.M. energy range 180–209 GeV, where  $E_\gamma > 2.5 \text{ GeV}$ , the photon has a polar angle  $|\cos\theta_\gamma| < 0.975$  and is well isolated from the nearest jet and charged lepton, and the effective masses of both fermion-antifermion systems agree with the  $W$  mass within  $3\Gamma_W$ . The measured differential cross section as a function of the photon energy and photon polar angle is used to extract the 95% CL limits:  $-0.020 \text{ GeV}^{-2} < a_0/\Lambda^2 < 0.020 \text{ GeV}^{-2}$ ,  $-0.053 \text{ GeV}^{-2} < a_c/\Lambda^2 < 0.037 \text{ GeV}^{-2}$  and  $-0.16 \text{ GeV}^{-2} < a_n/\Lambda^2 < 0.15 \text{ GeV}^{-2}$ .

<sup>8</sup> ABBIENDI 04L select  $20 e^+e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  acoplanar events in the energy range 180–209 GeV and  $176 e^+e^- \rightarrow q\bar{q}\gamma\gamma$  events in the energy range 130–209 GeV. These samples are used to constrain possible anomalous  $W^+W^-\gamma\gamma$  and  $ZZ\gamma\gamma$  quartic couplings. Further combining with the  $W^+W^-\gamma$  sample of ABBIENDI 04B the following one-parameter 95% CL limits are obtained:  $-0.007 < a_0^Z/\Lambda^2 < 0.023 \text{ GeV}^{-2}$ ,  $-0.029 < a_c^Z/\Lambda^2 < 0.029 \text{ GeV}^{-2}$ ,  $-0.020 < a_0^W/\Lambda^2 < 0.020 \text{ GeV}^{-2}$ ,  $-0.052 < a_c^W/\Lambda^2 < 0.037 \text{ GeV}^{-2}$ .

<sup>9</sup> In the CM energy range 183 to 209 GeV HEISTER 04A select  $30 e^+e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  events with two acoplanar, high energy and high transverse momentum photons. The photon-photon acoplanarity is required to be  $> 5^\circ$ ,  $E_\gamma/\sqrt{s} > 0.025$  (the more energetic photon having energy  $> 0.2\sqrt{s}$ ),  $p_{T_\gamma}/E_{\text{beam}} > 0.05$  and  $|\cos\theta_\gamma| < 0.94$ . A likelihood fit to the photon energy and recoil missing mass yields the following one-parameter 95% CL limits:  $-0.012 < a_0^Z/\Lambda^2 < 0.019 \text{ GeV}^{-2}$ ,  $-0.041 < a_c^Z/\Lambda^2 < 0.044 \text{ GeV}^{-2}$ ,  $-0.060 < a_0^W/\Lambda^2 < 0.055 \text{ GeV}^{-2}$ ,  $-0.099 < a_c^W/\Lambda^2 < 0.093 \text{ GeV}^{-2}$ .

<sup>10</sup> ABDALLAH 03i select  $122 e^+e^- \rightarrow W^+W^-\gamma$  events in the C.M. energy range 189–209 GeV, where  $E_\gamma > 5 \text{ GeV}$ , the photon has a polar angle  $|\cos\theta_\gamma| < 0.95$  and is well isolated from the nearest charged fermion. A fit to the photon energy spectrum yields  $a_c/\Lambda^2 = 0.000 \pm 0.019 \text{ GeV}^{-2}$ ,  $a_0/\Lambda^2 = -0.004 \pm 0.018 \text{ GeV}^{-2}$ ,  $\bar{a}_0/\Lambda^2 = -0.007 \pm 0.019 \text{ GeV}^{-2}$ ,  $a_n/\Lambda^2 = -0.09 \pm 0.16 \text{ GeV}^{-2}$ , and  $\bar{a}_n/\Lambda^2 = +0.05 \pm 0.07 \text{ GeV}^{-2}$ , keeping the other parameters fixed to their Standard Model values(0). The 95% CL limits are:  $-0.063 \text{ GeV}^{-2} < a_c/\Lambda^2 < +0.032 \text{ GeV}^{-2}$ ,  $-0.020 \text{ GeV}^{-2} < a_0/\Lambda^2 < +0.020 \text{ GeV}^{-2}$ ,  $-0.020 \text{ GeV}^{-2} < \bar{a}_0/\Lambda^2 < +0.020 \text{ GeV}^{-2}$ ,  $-0.18 \text{ GeV}^{-2} < a_n/\Lambda^2 < +0.14 \text{ GeV}^{-2}$ ,  $-0.16 \text{ GeV}^{-2} < \bar{a}_n/\Lambda^2 < +0.17 \text{ GeV}^{-2}$ .

<sup>11</sup> ACHARD 02F select  $86 e^+e^- \rightarrow W^+W^-\gamma$  events at 192–207 GeV, where  $E_\gamma > 5 \text{ GeV}$  and the photon is well isolated. They also select 43 acoplanar  $e^+e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  events in this energy range, where the photon energies are  $> 5 \text{ GeV}$  and  $> 1 \text{ GeV}$  and the photon polar angles are between  $14^\circ$  and  $166^\circ$ . All these 43 events are in the recoil mass region corresponding to the Z (75–110 GeV). Using the shape and normalization of the photon spectra in the  $W^+W^-\gamma$  events, and combining with the 42 event sample from 189 GeV data (ACCIARRI 00T), they obtain:  $a_0/\Lambda^2 = 0.000 \pm 0.010 \text{ GeV}^{-2}$ ,  $a_c/\Lambda^2 = -0.013 \pm 0.023 \text{ GeV}^{-2}$ , and  $a_n/\Lambda^2 = -0.002 \pm 0.076 \text{ GeV}^{-2}$ . Further combining the analyses of  $W^+W^-\gamma$  events with the low recoil mass region of  $\nu\bar{\nu}\gamma\gamma$  events (including samples collected at 183 + 189 GeV), they obtain the following one-parameter 95% CL limits:  $-0.015 \text{ GeV}^{-2} < a_0/\Lambda^2 < 0.015 \text{ GeV}^{-2}$ ,  $-0.048 \text{ GeV}^{-2} < a_c/\Lambda^2 < 0.026 \text{ GeV}^{-2}$ , and  $-0.14 \text{ GeV}^{-2} < a_n/\Lambda^2 < 0.13 \text{ GeV}^{-2}$ .

### W REFERENCES

PDG 16 Chin. Phys. C (PDG Collab.)

AAD 15N PRL 115 031802 (G. Aad et al., ATLAS Collab.)

KHACHATRYAN... 15D PRL 114 051801 (V. Khachatryan et al., CMS Collab.)

AAD 14AM PRL 113 141803 (G. Aad et al., ATLAS Collab.)

AAD 14Y JHEP 1404 031 (G. Aad et al., ATLAS Collab.)

AALTONEN 14D PR D89 072003 (T. Aaltonen et al., CDF Collab.)

ABAZOV 14N PR D89 012005 (V.M. Abazov et al., DO Collab.)

CHATRCHYAN 14AB PR D89 092005 (S. Chatrchyan et al., CMS Collab.)

CHATRCHYAN 14Q PR D90 032008 (S. Chatrchyan et al., CMS Collab.)

AAD 13AL PR D88 079006 (errat.) (G. Aad et al., ATLAS Collab.)

AAD 13AN PR D87 112003 (G. Aad et al., ATLAS Collab.)

Also PR D91 119901 (errat.) (G. Aad et al., ATLAS Collab.)

AALTONEN 13N PR D88 052018 (T. Aaltonen et al., CDF and DO Collabs.)

ABAZOV 13D PR D88 012005 (V.M. Abazov et al., DO Collab.)

CHATRCHYAN 13AA JHEP 1307 116 (S. Chatrchyan et al., CMS Collab.)

CHATRCHYAN 13BF EPJ C73 2610 (S. Chatrchyan et al., CMS Collab.)

SCHAEEL 13A PRPL 532 119 (S. Schaeel et al., (ALEPH Collab., DELPHI, L3+))

AAD 12AC PL B712 289 (G. Aad et al., ATLAS Collab.)

AAD 12BX PL B717 49 (G. Aad et al., ATLAS Collab.)

AAD 12CD EPJ C72 2173 (G. Aad et al., ATLAS Collab.)

AAD 12V PL B709 341 (G. Aad et al., ATLAS Collab.)

AALTONEN 12AC PR D86 031104 (T. Aaltonen et al., CDF Collab.)

AALTONEN 12E PRL 108 151803 (T. Aaltonen et al., CDF Collab.)

AALTONEN 12W PR D85 032001 (T. Aaltonen et al., CDF Collab.)

ABAZOV 12AG PL B718 451 (V.M. Abazov et al., DO Collab.)

ABAZOV 12F PRL 108 151804 (V.M. Abazov et al., DO Collab.)

ABAZOV 11 PL B695 67 (V.M. Abazov et al., DO Collab.)

ABAZOV 11AC PRL 107 241803 (V.M. Abazov et al., DO Collab.)

BLINOV 11 PL B699 287 (A.E. Blinov, A.S. Rudenko (NOVO))

CHATRCHYAN 11M PL B701 535 (S. Chatrchyan et al., CMS Collab.)

AALTONEN 10K PRL 104 201801 (T. Aaltonen et al., CDF Collab.)

Also PRL 105 019905 (errat.) (T. Aaltonen et al., (DELPHI Collab.))

ABDALLAH 10 EPJ C66 35 (J. Abdallah et al., (H1 Collab.))

AARON 09B EPJ C64 251 (F.D. Aaron et al., (DO Collab.))

ABAZOV 09AB PRL 103 141801 (V.M. Abazov et al., (DO Collab.))

ABAZOV 09AD PR D80 053012 (V.M. Abazov et al., (DO Collab.))

ABAZOV 09AJ PRL 103 191801 (V.M. Abazov et al., (DO Collab.))

ABAZOV 09AK PRL 103 231802 (V.M. Abazov et al., (DO Collab.))

AALTONEN 08B PRL 100 071801 (T. Aaltonen et al., (CDF Collab.))

ABAZOV 08R PRL 100 241805 (V.M. Abazov et al., (DO Collab.))

ABDALLAH 08A EPJ C55 1 (J. Abdallah et al., (DELPHI Collab.))

ABDALLAH 08C EPJ C54 345 (J. Abdallah et al., (DELPHI Collab.))

AALTONEN 07F PRL 99 151801 (T. Aaltonen et al., (CDF Collab.))

Also PR D77 112001 (T. Aaltonen et al., (CDF Collab.))

AALTONEN 07L PR D76 111103 (T. Aaltonen et al., (CDF Collab.))

ABAZOV 07Z PR D76 111104 (V.M. Abazov et al., (DO Collab.))

ABBIENDI 07A EPJ C52 767 (G. Abbiendi et al., (OPAL Collab.))

ABAZOV 06H PR D74 057101 (V.M. Abazov et al., (DO Collab.))

Also PR D74 059904 (errat.) (V.M. Abazov et al., (DO Collab.))

ABBIENDI 06 EPJ C45 307 (G. Abbiendi et al., (OPAL Collab.))

ABBIENDI 06A EPJ C45 291 (G. Abbiendi et al., (OPAL Collab.))

ACHARD 06 EPJ C45 569 (P. Achard et al., (L3 Collab.))

AKTAS 06 PL B632 35 (A. Aktas et al., (H1 Collab.))

SCHAEEL 06 EPJ C47 309 (S. Schaeel et al., (ALEPH Collab.))

ABAZOV 05J PR D71 091108 (V.M. Abazov et al., (DO Collab.))

ABAZOV 05S PRL 95 141802 (V.M. Abazov et al., (DO Collab.))

SCHAEEL 05A PL B614 7 (S. Schaeel et al., (ALEPH Collab.))

ABAZOV 04D PR D70 092008 (V.M. Abazov et al., (CDF Collab., DO Collab.))

ABBIENDI 04B PL B580 17 (G. Abbiendi et al., (OPAL Collab.))

ABBIENDI 04L EPJ C33 463 (G. Abbiendi et al., (OPAL Collab.))

ABBIENDI 04N PR D70 032005 (G. Abbiendi et al., (OPAL Collab.))

ABDALLAH 04G EPJ C34 127 (J. Abdallah et al., (DELPHI Collab.))

ACHARD 04D PL B586 151 (P. Achard et al., (L3 Collab.))

ACHARD 04J PL B600 22 (P. Achard et al., (L3 Collab.))

HEISTER 04A PL B602 31 (A. Heister et al., (ALEPH Collab.))

SCHAEEL 04A EPJ C38 147 (S. Schaeel et al., (ALEPH Collab.))

ABBIENDI 03C EPJ C26 321 (G. Abbiendi et al., (OPAL Collab.))

ABDALLAH 03I EPJ C31 139 (J. Abdallah et al., (DELPHI Collab.))

ABAZOV 02D PR D66 012001 (V.M. Abazov et al., (DO Collab.))

ABAZOV 02C PR D66 032008 (V.M. Abazov et al., (DO Collab.))

ACHARD 02F PL B527 29 (P. Achard et al., (L3 Collab.))

CHEKANOV 02C PL B539 197 (S. Chekanov et al., (ZEUS Collab.))

ABBIENDI 01H EPJ C19 229 (G. Abbiendi et al., (OPAL Collab.))

ABREU 01I PL B502 9 (P. Abreu et al., (DELPHI Collab.))

AFFOLDER 01E PR D64 052001 (T. Affolder et al., (CDF Collab.))

ABBIENDI 00V PL B490 71 (G. Abbiendi et al., (OPAL Collab.))

ABBOTT 00B PR D61 072001 (B. Abbott et al., (DO Collab.))

ABBOTT 00D PRL 84 5710 (B. Abbott et al., (DO Collab.))

ABREU.P 00F EPJ C18 203 (P. Abreu et al., (DELPHI Collab.))

Also EPJ C25 493 (errat.) (P. Abreu et al., (DELPHI Collab.))

ACCIARRI 00T PL B490 187 (M. Acciarri et al., (L3 Collab.))

AFFOLDER 00M PRL 85 3347 (T. Affolder et al., (CDF Collab.))

BREITWEG 00L PL B471 411 (J. Breitweg et al., (ZEUS Collab.))

BREITWEG 00D EPJ C12 411 (J. Breitweg et al., (ZEUS Collab.))

EBOLI 00M MPL A15 1 (O. Eboli, M. Gonzalez-Garcia, S. Novaes)

ABBIENDI 99N PL B453 153 (G. Abbiendi et al., (OPAL Collab.))

ABBOTT 99H PR D60 052003 (B. Abbott et al., (DO Collab.))

ABBOTT 99I PR D60 072002 (B. Abbott et al., (DO Collab.))

ABREU 99L PL B459 382 (P. Abreu et al., (DELPHI Collab.))

ACCIARRI 99P PL B454 386 (M. Acciarri et al., (L3 Collab.))

ACCIARRI 99Q PL B467 171 (M. Acciarri et al., (L3 Collab.))

BARATE 99I PL B453 107 (R. Barate et al., (ALEPH Collab.))

BARATE 99J PL B462 389 (R. Barate et al., (ALEPH Collab.))

BARATE 99M PL B465 349 (R. Barate et al., (ALEPH Collab.))

ABBOTT 98N PR D58 092003 (B. Abbott et al., (DO Collab.))

ABBOTT 98P PR D58 012002 (B. Abbott et al., (DO Collab.))

ABE 98H PR D58 031101 (F. Abe et al., (CDF Collab.))

ABE 98P PR D58 091101 (F. Abe et al., (CDF Collab.))

ABREU 98C PL B416 233 (P. Abreu et al., (DELPHI Collab.))

ABREU 98N PL B439 209 (P. Abreu et al., (DELPHI Collab.))

BARATE 97 PL B401 347 (R. Barate et al., (ALEPH Collab.))

BARATE 97S PL B415 435 (R. Barate et al., (ALEPH Collab.))

ABACHI 95D PRL 75 1456 (S. Abachi et al., (DO Collab.))

ABE 95C PRL 74 341 (F. Abe et al., (CDF Collab.))

ABE 95G PRL 74 1936 (F. Abe et al., (CDF Collab.))

ABE 95P PRL 75 11 (F. Abe et al., (CDF Collab.))

Also PR D52 4784 (F. Abe et al., (CDF Collab.))

ABE 95W PR D52 2624 (F. Abe et al., (CDF Collab.))

Also PRL 73 220 (F. Abe et al., (CDF Collab.))

ABE 92E PRL 68 3398 (F. Abe et al., (CDF Collab.))

ABE 92I PRL 69 28 (F. Abe et al., (CDF Collab.))

ALITTI 92B PL B276 365 (J. Alitti et al., (UA2 Collab.))

ALITTI 92B PL B276 354 (J. Alitti et al., (UA2 Collab.))

ALITTI 92C PL B277 194 (J. Alitti et al., (UA2 Collab.))

ALITTI 92D PL B277 203 (J. Alitti et al., (UA2 Collab.))

ALITTI 92F PL B280 137 (J. Alitti et al., (UA2 Collab.))

SAMUEL 92 PL B280 124 (M.A. Samuel et al., (OKSU, CARL))

ABE 91C PR D44 29 (F. Abe et al., (CDF Collab.))

ALBAJAR 91 PL B253 503 (C. Albajar et al., (UA1 Collab.))

ALITTI 91C ZPHY C52 209 (J. Alitti et al., (UA2 Collab.))

SAMUEL 91 PRL 67 9 (M.A. Samuel et al., (OKSU, CARL))

Also PRL 67 2920 (erratum) (M.A. Samuel et al., (OKSU, CARL))

ABE 90G PRL 65 2243 (F. Abe et al., (CDF Collab.))

Also PR D43 2070 (F. Abe et al., (CDF Collab.))

ALBAJAR 90 PL B241 283 (C. Albajar et al., (UA1 Collab.))

ALITTI 90B PL B241 150 (J. Alitti et al., (UA2 Collab.))

ABE 89I PRL 62 1005 (F. Abe et al., (CDF Collab.))

ALBAJAR 89 ZPHY C44 15 (C. Albajar et al., (UA1 Collab.))

BAUR 88 NP B308 127 (U. Baur, D. Zeppenfeld (FSU, WISC))

GRIFOLS 88 JUMP A3 225 (J.A. Grifols, S. Peris, J. Sola (BARC, DESY))

Also PL B197 437 (J.A. Grifols, S. Peris, J. Sola (BARC, DESY))

ALBAJAR 87 PL B185 233 (C. Albajar et al., (UA1 Collab.))

ANSARI 87 PL B186 440 (R. Ansari et al., (UA2 Collab.))

GROTCH 87 PR D36 2153 (H. Grotch, R.W. Robinett (FSU))

HAGIWARA 87 NP B282 253 (K. Hagiwara et al., (KEK, UCLA, FSU))

VANDERBIJ 87 PR D35 1088 (J.J. van der Bij (FNAL))

GRAU 85 PL 154B 283 (A. Grau, J.A. Grifols (BARC))

SUZUKI 85 PL 153B 289 (M. Suzuki (LBL))

ARNISON 84D PL 134B 469 (G.T.J. Armonson et al., (UA1 Collab.))

HERZOG 84 PL 148B 355 (F. Herzog (WISC))

Also PL 155B 468 (erratum) (F. Herzog (WISC))

ARNISON 83 PL 122B 103 (G.T.J. Armonson et al., (UA1 Collab.))

BANNER 83B PL 122B 476 (M. Banner et al., (UA2 Collab.))



$J = 1$

### THE Z BOSON

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Precision measurements at the Z-boson resonance using electron-positron colliding beams began in 1989 at the SLC and at LEP. During 1989–95, the four LEP experiments (ALEPH, DELPHI, L3, OPAL) made high-statistics studies of the production and decay properties of the Z. Although the SLD experiment at the SLC collected much lower statistics, it was able to match the precision of LEP experiments in determining



the effective electroweak mixing angle  $\sin^2\bar{\theta}_W$  and the rates of  $Z$  decay to  $b$ - and  $c$ -quarks, owing to availability of polarized electron beams, small beam size, and stable beam spot.

The  $Z$ -boson properties reported in this section may broadly be categorized as:

- The standard ‘lineshape’ parameters of the  $Z$  consisting of its mass,  $M_Z$ , its total width,  $\Gamma_Z$ , and its partial decay widths,  $\Gamma(\text{hadrons})$ , and  $\Gamma(\ell\bar{\ell})$  where  $\ell = e, \mu, \tau, \nu$ ;
- $Z$  asymmetries in leptonic decays and extraction of  $Z$  couplings to charged and neutral leptons;
- The  $b$ - and  $c$ -quark-related partial widths and charge asymmetries which require special techniques;
- Determination of  $Z$  decay modes and the search for modes that violate known conservation laws;
- Average particle multiplicities in hadronic  $Z$  decay;
- $Z$  anomalous couplings.

The effective vector and axial-vector coupling constants describing the  $Z$ -to-fermion coupling are also measured in  $p\bar{p}$  and  $ep$  collisions at the Tevatron and at HERA. The corresponding cross-section formulae are given in Section 39 (Cross-section formulae for specific processes) and Section 16 (Structure Functions) in this *Review*. In this minireview, we concentrate on the measurements in  $e^+e^-$  collisions at LEP and SLC.

The standard ‘lineshape’ parameters of the  $Z$  are determined from an analysis of the production cross sections of these final states in  $e^+e^-$  collisions. The  $Z \rightarrow \nu\bar{\nu}(\gamma)$  state is identified directly by detecting single photon production and indirectly by subtracting the visible partial widths from the total width. Inclusion in this analysis of the forward-backward asymmetry of charged leptons,  $A_{FB}^{(0,\ell)}$ , of the  $\tau$  polarization,  $P(\tau)$ , and its forward-backward asymmetry,  $P(\tau)^{fb}$ , enables the separate determination of the effective vector ( $\bar{g}_V$ ) and axial vector ( $\bar{g}_A$ ) couplings of the  $Z$  to these leptons and the ratio ( $\bar{g}_V/\bar{g}_A$ ), which is related to the effective electroweak mixing angle  $\sin^2\bar{\theta}_W$  (see the “Electroweak Model and Constraints on New Physics” review).

Determination of the  $b$ - and  $c$ -quark-related partial widths and charge asymmetries involves tagging the  $b$  and  $c$  quarks for which various methods are employed: requiring the presence of a high momentum prompt lepton in the event with high transverse momentum with respect to the accompanying jet; impact parameter and lifetime tagging using precision vertex measurement with high-resolution detectors; application of neural-network techniques to classify events as  $b$  or non- $b$  on a statistical basis using event-shape variables; and using the presence of a charmed meson ( $D/D^*$ ) or a kaon as a tag.

### ***Z*-parameter determination**

LEP was run at energy points on and around the  $Z$  mass (88–94 GeV) constituting an energy ‘scan.’ The shape of the cross-section variation around the  $Z$  peak can be described by a Breit-Wigner *ansatz* with an energy-dependent

total width [1–3]. The **three** main properties of this distribution, viz., the **position** of the peak, the **width** of the distribution, and the **height** of the peak, determine respectively the values of  $M_Z$ ,  $\Gamma_Z$ , and  $\Gamma(e^+e^-) \times \Gamma(f\bar{f})$ , where  $\Gamma(e^+e^-)$  and  $\Gamma(f\bar{f})$  are the electron and fermion partial widths of the  $Z$ . The quantitative determination of these parameters is done by writing analytic expressions for these cross sections in terms of the parameters, and fitting the calculated cross sections to the measured ones by varying these parameters, taking properly into account all the errors. Single-photon exchange ( $\sigma_\gamma^0$ ) and  $\gamma$ - $Z$  interference ( $\sigma_{\gamma Z}^0$ ) are included, and the large ( $\sim 25\%$ ) initial-state radiation (ISR) effects are taken into account by convoluting the analytic expressions over a ‘Radiator Function’ [1–5]  $H(s, s')$ . Thus for the process  $e^+e^- \rightarrow f\bar{f}$ :

$$\sigma_f(s) = \int H(s, s') \sigma_f^0(s') ds' \quad (1)$$

$$\sigma_f^0(s) = \sigma_Z^0 + \sigma_\gamma^0 + \sigma_{\gamma Z}^0 \quad (2)$$

$$\sigma_Z^0 = \frac{12\pi}{M_Z^2} \frac{\Gamma(e^+e^-)\Gamma(f\bar{f})}{\Gamma_Z^2} \frac{s \Gamma_Z^2}{(s - M_Z^2)^2 + s^2 \Gamma_Z^2 / M_Z^2} \quad (3)$$

$$\sigma_\gamma^0 = \frac{4\pi\alpha^2(s)}{3s} Q_f^2 N_c^f \quad (4)$$

$$\begin{aligned} \sigma_{\gamma Z}^0 = & - \frac{2\sqrt{2}\alpha(s)}{3} (Q_f G_F N_c^f \mathcal{G}_V^e \mathcal{G}_V^f) \\ & \times \frac{(s - M_Z^2) M_Z^2}{(s - M_Z^2)^2 + s^2 \Gamma_Z^2 / M_Z^2} \end{aligned} \quad (5)$$

where  $Q_f$  is the charge of the fermion,  $N_c^f = 3$  for quarks and 1 for leptons, and  $\mathcal{G}_V^f$  is the vector coupling of the  $Z$  to the fermion-antifermion pair  $f\bar{f}$ .

Since  $\sigma_{\gamma Z}^0$  is expected to be much less than  $\sigma_Z^0$ , the LEP Collaborations have generally calculated the interference term in the framework of the Standard Model. This fixing of  $\sigma_{\gamma Z}^0$  leads to a tighter constraint on  $M_Z$ , and consequently a smaller error on its fitted value. It is possible to relax this constraint and carry out the fit within the S-matrix framework, which is briefly described in the next section.

In the above framework, the QED radiative corrections have been explicitly taken into account by convoluting over the ISR and allowing the electromagnetic coupling constant to run [6]:  $\alpha(s) = \alpha/(1 - \Delta\alpha)$ . On the other hand, weak radiative corrections that depend upon the assumptions of the electroweak theory and on the values of  $M_{\text{top}}$  and  $M_{\text{Higgs}}$  are accounted for by **absorbing them into the couplings**, which are then called the *effective* couplings  $\mathcal{G}_V$  and  $\mathcal{G}_A$  (or alternatively the effective parameters of the  $\star$  scheme of Kennedy and Lynn [7].)

$\mathcal{G}_V^f$  and  $\mathcal{G}_A^f$  are complex numbers with small imaginary parts. As experimental data does not allow simultaneous extraction of both real and imaginary parts of the effective couplings, the convention  $g_A^f = \text{Re}(\mathcal{G}_A^f)$  and  $g_V^f = \text{Re}(\mathcal{G}_V^f)$  is used and the imaginary parts are added in the fitting code [4].

Defining

$$A_f = 2 \frac{g_V^f \cdot g_A^f}{(g_V^f)^2 + (g_A^f)^2} \quad (6)$$

# Gauge & Higgs Boson Particle Listings

## Z

the lowest-order expressions for the various lepton-related asymmetries on the  $Z$  pole are [8–10]  $A_{FB}^{(0,\ell)} = (3/4)A_e A_f$ ,  $P(\tau) = -A_\tau$ ,  $P(\tau)^{fb} = -(3/4)A_e$ ,  $A_{LR} = A_e$ . The full analysis takes into account the energy-dependence of the asymmetries. Experimentally  $A_{LR}$  is defined as  $(\sigma_L - \sigma_R)/(\sigma_L + \sigma_R)$ , where  $\sigma_{L(R)}$  are the  $e^+e^- \rightarrow Z$  production cross sections with left- (right)-handed electrons.

The definition of the partial decay width of the  $Z$  to  $f\bar{f}$  includes the effects of QED and QCD final-state corrections, as well as the contribution due to the imaginary parts of the couplings:

$$\Gamma(f\bar{f}) = \frac{G_F M_Z^3}{6\sqrt{2}\pi} N_c^f \left( |g_A^f|^2 R_A^f + |g_V^f|^2 R_V^f \right) + \Delta_{ew/QCD} \quad (7)$$

where  $R_V^f$  and  $R_A^f$  are radiator factors to account for final state QED and QCD corrections, as well as effects due to nonzero fermion masses, and  $\Delta_{ew/QCD}$  represents the non-factorizable electroweak/QCD corrections.

### S-matrix approach to the Z

While most experimental analyses of LEP/SLC data have followed the ‘Breit-Wigner’ approach, an alternative S-matrix-based analysis is also possible. The  $Z$ , like all unstable particles, is associated with a complex pole in the S matrix. The pole position is process-independent and gauge-invariant. The mass,  $\bar{M}_Z$ , and width,  $\bar{\Gamma}_Z$ , can be defined in terms of the pole in the energy plane via [11–14]

$$\bar{s} = \bar{M}_Z^2 - i\bar{M}_Z\bar{\Gamma}_Z \quad (8)$$

leading to the relations

$$\begin{aligned} \bar{M}_Z &= M_Z / \sqrt{1 + \Gamma_Z^2/M_Z^2} \\ &\approx M_Z - 34.1 \text{ MeV} \end{aligned} \quad (9)$$

$$\begin{aligned} \bar{\Gamma}_Z &= \Gamma_Z / \sqrt{1 + \Gamma_Z^2/M_Z^2} \\ &\approx \Gamma_Z - 0.9 \text{ MeV} . \end{aligned} \quad (10)$$

The LEP collaborations [15] have analyzed their data using the S-matrix approach as defined in Eq. (8), in addition to the conventional one. They observe a downward shift in the  $Z$  mass as expected.

### Handling the large-angle $e^+e^-$ final state

Unlike other  $f\bar{f}$  decay final states of the  $Z$ , the  $e^+e^-$  final state has a contribution not only from the  $s$ -channel but also from the  $t$ -channel and  $s$ - $t$  interference. The full amplitude is not amenable to fast calculation, which is essential if one has to carry out minimization fits within reasonable computer time. The usual procedure is to calculate the non- $s$  channel part of the cross section separately using the Standard Model programs ALIBABA [16] or TOPAZ0 [17], with the measured value of  $M_{\text{top}}$ , and  $M_{\text{Higgs}} = 150 \text{ GeV}$ , and add it to the  $s$ -channel cross section calculated as for other channels. This leads to two additional sources of error in the analysis: firstly, the theoretical calculation in ALIBABA itself is known to be

accurate to  $\sim 0.5\%$ , and secondly, there is uncertainty due to the error on  $M_{\text{top}}$  and the unknown value of  $M_{\text{Higgs}}$  (100–1000 GeV). These errors are propagated into the analysis by including them in the systematic error on the  $e^+e^-$  final state. As these errors are common to the four LEP experiments, this is taken into account when performing the LEP average.

### Errors due to uncertainty in LEP energy determination [18–23]

The systematic errors related to the LEP energy measurement can be classified as:

- The absolute energy scale error;
- Energy-point-to-energy-point errors due to the non-linear response of the magnets to the exciting currents;
- Energy-point-to-energy-point errors due to possible higher-order effects in the relationship between the dipole field and beam energy;
- Energy reproducibility errors due to various unknown uncertainties in temperatures, tidal effects, corrector settings, RF status, *etc.*

Precise energy calibration was done outside normal data-taking using the resonant depolarization technique. Run-time energies were determined every 10 minutes by measuring the relevant machine parameters and using a model which takes into account all the known effects, including leakage currents produced by trains in the Geneva area and the tidal effects due to gravitational forces of the Sun and the Moon. The LEP Energy Working Group has provided a covariance matrix from the determination of LEP energies for the different running periods during 1993–1995 [18].

### Choice of fit parameters

The LEP Collaborations have chosen the following primary set of parameters for fitting:  $M_Z$ ,  $\Gamma_Z$ ,  $\sigma_{\text{hadron}}^0$ ,  $R(\text{lepton})$ ,  $A_{FB}^{(0,\ell)}$ , where  $R(\text{lepton}) = \Gamma(\text{hadrons})/\Gamma(\text{lepton})$ ,  $\sigma_{\text{hadron}}^0 = 12\pi\Gamma(e^+e^-)\Gamma(\text{hadrons})/M_Z^2\Gamma_Z^2$ . With a knowledge of these fitted parameters and their covariance matrix, any other parameter can be derived. The main advantage of these parameters is that they form a physics motivated set of parameters with much reduced correlations.

Thus, the most general fit carried out to cross section and asymmetry data determines the **nine parameters**:  $M_Z$ ,  $\Gamma_Z$ ,  $\sigma_{\text{hadron}}^0$ ,  $R(e)$ ,  $R(\mu)$ ,  $R(\tau)$ ,  $A_{FB}^{(0,e)}$ ,  $A_{FB}^{(0,\mu)}$ ,  $A_{FB}^{(0,\tau)}$ . Assumption of lepton universality leads to a **five-parameter fit** determining  $M_Z$ ,  $\Gamma_Z$ ,  $\sigma_{\text{hadron}}^0$ ,  $R(\text{lepton})$ ,  $A_{FB}^{(0,\ell)}$ .

### Combining results from LEP and SLC experiments

With a steady increase in statistics over the years and improved understanding of the common systematic errors between LEP experiments, the procedures for combining results have evolved continuously [24]. The Line Shape Sub-group of the LEP Electroweak Working Group investigated the effects of these common errors, and devised a combination procedure for the precise determination of the  $Z$  parameters from LEP experiments. Using these procedures, this note also gives the

results after combining the final parameter sets from the four experiments, and these are the results quoted as the fit results in the  $Z$  listings below. Transformation of variables leads to values of derived parameters like partial decay widths and branching ratios to hadrons and leptons. Finally, transforming the LEP combined nine parameter set to  $(M_Z, \Gamma_Z, \sigma_{\text{hadron}}^0, g_A^f, g_V^f, f = e, \mu, \tau)$  using the average values of lepton asymmetry parameters  $(A_e, A_\mu, A_\tau)$  as constraints, leads to the best fitted values of the vector and axial-vector couplings  $(g_V, g_A)$  of the charged leptons to the  $Z$ .

Brief remarks on the handling of common errors and their magnitudes are given below. The identified common errors are those coming from

- (a) LEP energy-calibration uncertainties, and
- (b) the theoretical uncertainties in (i) the luminosity determination using small angle Bhabha scattering, (ii) estimating the non- $s$  channel contribution to large angle Bhabha scattering, (iii) the calculation of QED radiative effects, and (iv) the parametrization of the cross section in terms of the parameter set used.

#### Common LEP energy errors

All the collaborations incorporate in their fit the full LEP energy error matrix as provided by the LEP energy group for their intersection region [18]. The effect of these errors is separated out from that of other errors by carrying out fits with energy errors scaled up and down by  $\sim 10\%$  and redoing the fits. From the observed changes in the overall error matrix, the covariance matrix of the common energy errors is determined. Common LEP energy errors lead to uncertainties on  $M_Z$ ,  $\Gamma_Z$ , and  $\sigma_{\text{hadron}}^0$  of 1.7, 1.2 MeV, and 0.011 nb, respectively.

#### Common luminosity errors

BHLUMI 4.04 [25] is used by all LEP collaborations for small-angle Bhabha scattering leading to a common uncertainty in their measured cross sections of 0.061% [26]. BHLUMI does not include a correction for production of light fermion pairs. OPAL explicitly corrects for this effect and reduces their luminosity uncertainty to 0.054%, which is taken fully correlated with the other experiments. The other three experiments among themselves have a common uncertainty of 0.061%.

#### Common non- $s$ channel uncertainties

The same standard model programs ALIBABA [16] and TOPAZ0 [17] are used to calculate the non- $s$  channel contribution to the large angle Bhabha scattering [27]. As this contribution is a function of the  $Z$  mass, which itself is a variable in the fit, it is parametrized as a function of  $M_Z$  by each collaboration to properly track this contribution as  $M_Z$  varies in the fit. The common errors on  $R_e$  and  $A_{FB}^{(0,e)}$  are 0.024 and 0.0014 respectively, and are correlated between them.

#### Common theoretical uncertainties: QED

There are large initial-state photon and fermion pair radiation effects near the  $Z$  resonance, for which the best currently available evaluations include contributions up to  $\mathcal{O}(\alpha^3)$ . To

estimate the remaining uncertainties, different schemes are incorporated in the standard model programs ZFITTER [5], TOPAZ0 [17], and MIZA [28]. Comparing the different options leads to error estimates of 0.3 and 0.2 MeV on  $M_Z$  and  $\Gamma_Z$  respectively, and of 0.02% on  $\sigma_{\text{hadron}}^0$ .

#### Common theoretical uncertainties: parametrization of lineshape and asymmetries

To estimate uncertainties arising from ambiguities in the model-independent parametrization of the differential cross-section near the  $Z$  resonance, results from TOPAZ0 and ZFITTER were compared by using ZFITTER to fit the cross sections and asymmetries calculated using TOPAZ0. The resulting uncertainties on  $M_Z$ ,  $\Gamma_Z$ ,  $\sigma_{\text{hadron}}^0$ ,  $R(\text{lepton})$ , and  $A_{FB}^{(0,\ell)}$  are 0.1 MeV, 0.1 MeV, 0.001 nb, 0.004, and 0.0001 respectively.

Thus, the overall theoretical errors on  $M_Z$ ,  $\Gamma_Z$ ,  $\sigma_{\text{hadron}}^0$  are 0.3 MeV, 0.2 MeV, and 0.008 nb respectively; on each  $R(\text{lepton})$  is 0.004 and on each  $A_{FB}^{(0,\ell)}$  is 0.0001. Within the set of three  $R(\text{lepton})$ 's and the set of three  $A_{FB}^{(0,\ell)}$ 's, the respective errors are fully correlated.

All the theory-related errors mentioned above utilize Standard Model programs which need the Higgs mass and running electromagnetic coupling constant as inputs; uncertainties on these inputs will also lead to common errors. All LEP collaborations used the same set of inputs for Standard Model calculations:  $M_Z = 91.187$  GeV, the Fermi constant  $G_F = (1.16637 \pm 0.00001) \times 10^{-5} \text{ GeV}^{-2}$  [29],  $\alpha^{(5)}(M_Z) = 1/128.877 \pm 0.090$  [30],  $\alpha_s(M_Z) = 0.119$  [31],  $M_{\text{top}} = 174.3 \pm 5.1$  GeV [31] and  $M_{\text{Higgs}} = 150$  GeV. The only observable effect, on  $M_Z$ , is due to the variation of  $M_{\text{Higgs}}$  between 100–1000 GeV (due to the variation of the  $\gamma/Z$  interference term which is taken from the Standard Model):  $M_Z$  changes by +0.23 MeV per unit change in  $\log_{10} M_{\text{Higgs}}/\text{GeV}$ , which is not an error but a correction to be applied once  $M_{\text{Higgs}}$  is determined. The effect is much smaller than the error on  $M_Z$  ( $\pm 2.1$  MeV).

#### Methodology of combining the LEP experimental results

The LEP experimental results actually used for combination are slightly modified from those published by the experiments (which are given in the Listings below). This has been done in order to facilitate the procedure by making the inputs more consistent. These modified results are given explicitly in [24]. The main differences compared to the published results are (a) consistent use of ZFITTER 6.23 and TOPAZ0 (the published ALEPH results used ZFITTER 6.10); (b) use of the combined energy-error matrix, which makes a difference of 0.1 MeV on the  $M_Z$  and  $\Gamma_Z$  for L3 only as at that intersection the RF modeling uncertainties are the largest.

Thus, nine-parameter sets from all four experiments with their covariance matrices are used together with all the common errors correlations. A grand covariance matrix,  $V$ , is constructed and a combined nine-parameter set is obtained by minimizing  $\chi^2 = \Delta^T V^{-1} \Delta$ , where  $\Delta$  is the vector of residuals of the combined parameter set to the results of individual

# Gauge & Higgs Boson Particle Listings

## Z

experiments. Imposing lepton universality in the combination results in the combined five parameter set.

### Study of $Z \rightarrow b\bar{b}$ and $Z \rightarrow c\bar{c}$

In the sector of  $c$ - and  $b$ -physics, the LEP experiments have measured the ratios of partial widths  $R_b = \Gamma(Z \rightarrow b\bar{b})/\Gamma(Z \rightarrow \text{hadrons})$ , and  $R_c = \Gamma(Z \rightarrow c\bar{c})/\Gamma(Z \rightarrow \text{hadrons})$ , and the forward-backward (charge) asymmetries  $A_{FB}^{b\bar{b}}$  and  $A_{FB}^{c\bar{c}}$ . The SLD experiment at SLC has measured the ratios  $R_c$  and  $R_b$  and, utilizing the polarization of the electron beam, was able to obtain the final state coupling parameters  $A_b$  and  $A_c$  from a measurement of the left-right forward-backward asymmetry of  $b$ - and  $c$ -quarks. The high precision measurement of  $R_c$  at SLD was made possible owing to the small beam size and very stable beam spot at SLC, coupled with a highly precise CCD pixel detector. Several of the analyses have also determined other quantities, in particular the semileptonic branching ratios,  $B(b \rightarrow \ell^-)$ ,  $B(b \rightarrow c \rightarrow \ell^+)$ , and  $B(c \rightarrow \ell^+)$ , the average time-integrated  $B^0\bar{B}^0$  mixing parameter  $\bar{\chi}$  and the probabilities for a  $c$ -quark to fragment into a  $D^+$ , a  $D_s$ , a  $D^{*+}$ , or a charmed baryon. The latter measurements do not concern properties of the  $Z$  boson, and hence they do not appear in the Listing below. However, for completeness, we will report at the end of this minireview their values as obtained fitting the data contained in the  $Z$  section. All these quantities are correlated with the electroweak parameters, and since the mixture of  $b$  hadrons is different from the one at the  $\Upsilon(4S)$ , their values might differ from those measured at the  $\Upsilon(4S)$ .

All the above quantities are correlated to each other since:

- Several analyses (for example the lepton fits) determine more than one parameter simultaneously;
- Some of the electroweak parameters depend explicitly on the values of other parameters (for example  $R_b$  depends on  $R_c$ );
- Common tagging and analysis techniques produce common systematic uncertainties.

The LEP Electroweak Heavy Flavour Working Group has developed [32] a procedure for combining the measurements taking into account known sources of correlation. The combining procedure determines fourteen parameters: the six parameters of interest in the electroweak sector,  $R_b$ ,  $R_c$ ,  $A_{FB}^{b\bar{b}}$ ,  $A_{FB}^{c\bar{c}}$ ,  $A_b$  and  $A_c$  and, in addition,  $B(b \rightarrow \ell^-)$ ,  $B(b \rightarrow c \rightarrow \ell^+)$ ,  $B(c \rightarrow \ell^+)$ ,  $\bar{\chi}$ ,  $f(D^+)$ ,  $f(D_s)$ ,  $f(c_{\text{baryon}})$  and  $P(c \rightarrow D^{*+}) \times B(D^{*+} \rightarrow \pi^+ D^0)$ , to take into account their correlations with the electroweak parameters. Before the fit both the peak and off-peak asymmetries are translated to the common energy  $\sqrt{s} = 91.26$  GeV using the predicted energy-dependence from ZFITTER [5].

### Summary of the measurements and of the various kinds of analysis

The measurements of  $R_b$  and  $R_c$  fall into two classes. In the first, named single-tag measurement, a method for selecting  $b$  and  $c$  events is applied and the number of tagged events is counted. A second technique, named double-tag measurement,

has the advantage that the tagging efficiency is directly derived from the data thereby reducing the systematic error on the measurement.

The measurements in the  $b$ - and  $c$ -sector can be essentially grouped in the following categories:

- Lifetime (and lepton) double-tagging measurements of  $R_b$ . These are the most precise measurements of  $R_b$  and obviously dominate the combined result. The main sources of systematics come from the charm contamination and from estimating the hemisphere  $b$ -tagging efficiency correlation;
- Analyses with  $D/D^{*\pm}$  to measure  $R_c$ . These measurements make use of several different tagging techniques (inclusive/exclusive double tag, exclusive double tag, reconstruction of all weakly decaying charmed states) and no assumptions are made on the energy-dependence of charm fragmentation;
- A measurement of  $R_c$  using single leptons and assuming  $B(b \rightarrow c \rightarrow \ell^+)$ ;
- Lepton fits which use hadronic events with one or more leptons in the final state to measure the asymmetries  $A_{FB}^{b\bar{b}}$  and  $A_{FB}^{c\bar{c}}$ . Each analysis usually gives several other electroweak parameters. The dominant sources of systematics are due to lepton identification, to other semileptonic branching ratios and to the modeling of the semileptonic decay;
- Measurements of  $A_{FB}^{b\bar{b}}$  using lifetime tagged events with a hemisphere charge measurement. These measurements dominate the combined result;
- Analyses with  $D/D^{*\pm}$  to measure  $A_{FB}^{c\bar{c}}$  or simultaneously  $A_{FB}^{b\bar{b}}$  and  $A_{FB}^{c\bar{c}}$ ;
- Measurements of  $A_b$  and  $A_c$  from SLD, using several tagging methods (lepton, kaon,  $D/D^*$ , and vertex mass). These quantities are directly extracted from a measurement of the left-right forward-backward asymmetry in  $c\bar{c}$  and  $b\bar{b}$  production using a polarized electron beam.

### Averaging procedure

All the measurements are provided by the LEP and SLD Collaborations in the form of tables with a detailed breakdown of the systematic errors of each measurement and its dependence on other electroweak parameters.

The averaging proceeds via the following steps:

- Define and propagate a consistent set of external inputs such as branching ratios, hadron lifetimes, fragmentation models *etc.* All the measurements are checked to ensure that all use a common set of assumptions (for instance, since the QCD corrections for the forward-backward asymmetries are strongly dependent on the experimental conditions, the data are corrected before combining);

- Form the full (statistical and systematic) covariance matrix of the measurements. The systematic correlations between different analyses are calculated from the detailed error breakdown in the measurement tables. The correlations relating several measurements made by the same analysis are also used;
- Take into account any explicit dependence of a measurement on the other electroweak parameters. As an example of this dependence, we illustrate the case of the double-tag measurement of  $R_b$ , where  $c$ -quarks constitute the main background. The normalization of the charm contribution is not usually fixed by the data and the measurement of  $R_b$  depends on the assumed value of  $R_c$ , which can be written as:

$$R_b = R_b^{\text{meas}} + a(R_c) \frac{(R_c - R_c^{\text{used}})}{R_c}, \quad (11)$$

where  $R_b^{\text{meas}}$  is the result of the analysis which assumed a value of  $R_c = R_c^{\text{used}}$  and  $a(R_c)$  is the constant which gives the dependence on  $R_c$ ;

- Perform a  $\chi^2$  minimization with respect to the combined electroweak parameters.

After the fit the average peak asymmetries  $A_{FB}^{c\bar{c}}$  and  $A_{FB}^{b\bar{b}}$  are corrected for the energy shift from 91.26 GeV to  $M_Z$  and for QED (initial state radiation),  $\gamma$  exchange, and  $\gamma Z$  interference effects, to obtain the corresponding pole asymmetries  $A_{FB}^{0,c}$  and  $A_{FB}^{0,b}$ .

This averaging procedure, using the fourteen parameters described above, and applied to the data contained in the Z particle listing below, gives the following results (where the last 8 parameters do not depend directly on the Z):

$$R_b^0 = 0.21629 \pm 0.00066$$

$$R_c^0 = 0.1721 \pm 0.0030$$

$$A_{FB}^{0,b} = 0.0992 \pm 0.0016$$

$$A_{FB}^{0,c} = 0.0707 \pm 0.0035$$

$$A_b = 0.923 \pm 0.020$$

$$A_c = 0.670 \pm 0.027$$

$$B(b \rightarrow \ell^-) = 0.1071 \pm 0.0022$$

$$B(b \rightarrow c \rightarrow \ell^+) = 0.0801 \pm 0.0018$$

$$B(c \rightarrow \ell^+) = 0.0969 \pm 0.0031$$

$$\bar{\chi} = 0.1250 \pm 0.0039$$

$$f(D^+) = 0.235 \pm 0.016$$

$$f(D_s) = 0.126 \pm 0.026$$

$$f(c_{\text{baryon}}) = 0.093 \pm 0.022$$

$$P(c \rightarrow D^{*+}) \times B(D^{*+} \rightarrow \pi^+ D^0) = 0.1622 \pm 0.0048$$

Among the non-electroweak observables, the B semileptonic branching fraction  $B(b \rightarrow \ell^-)$  is of special interest, since the dominant error source on this quantity is the dependence on the semileptonic decay model for  $b \rightarrow \ell^-$ , with  $\Delta B(b \rightarrow \ell^-)_{b \rightarrow \ell^- \text{ model}} = 0.0012$ . Extensive studies have been made to understand the size of this error. Among the electroweak quantities, the quark asymmetries with leptons depend also on the semileptonic decay model, while the asymmetries using other methods usually do not. The fit implicitly requires that the different methods give consistent results and this effectively constrains the decay model, and thus reduces in principle the error from this source in the fit result.

To obtain a conservative estimate of the modelling error, the above fit has been repeated removing all asymmetry measurements. The results of the fit on B-decay related observables are [24]:  $B(b \rightarrow \ell^-) = 0.1069 \pm 0.0022$ , with  $\Delta B(b \rightarrow \ell^-)_{b \rightarrow \ell^- \text{ model}} = 0.0013$ ,  $B(b \rightarrow c \rightarrow \ell^+) = 0.0802 \pm 0.0019$  and  $\bar{\chi} = 0.1259 \pm 0.0042$ .

## References

1. R.N. Cahn, Phys. Rev. **D36**, 2666 (1987).
2. F.A. Berends *et al.*, “Z Physics at LEP 1,” CERN Report 89-08 (1989), Vol. 1, eds. G. Altarelli, R. Kleiss, and C. Verzegnassi, p. 89.
3. A. Borrelli *et al.*, Nucl. Phys. **B333**, 357 (1990).
4. D. Bardin and G. Passarino, “Upgrading of Precision Calculations for Electroweak Observables,” **hep-ph/9803425**; D. Bardin, G. Passarino, and M. Grünewald, “Precision Calculation Project Report,” **hep-ph/9902452**.
5. D. Bardin *et al.*, Z. Phys. **C44**, 493 (1989); Comp. Phys. Comm. **59**, 303 (1990); D. Bardin *et al.*, Nucl. Phys. **B351**, 1 (1991); Phys. Lett. **B255**, 290 (1991), and CERN-TH/6443/92 (1992); Comp. Phys. Comm. **133**, 229 (2001).
6. G. Burgers *et al.*, “Z Physics at LEP 1,” CERN Report 89-08 (1989), Vol. 1, eds. G. Altarelli, R. Kleiss, and C. Verzegnassi, p. 55.
7. D.C. Kennedy and B.W. Lynn, Nucl. Phys. **B322**, 1 (1989).
8. M. Consoli *et al.*, “Z Physics at LEP 1,” CERN Report 89-08 (1989), Vol. 1, eds. G. Altarelli, R. Kleiss, and C. Verzegnassi, p. 7.
9. M. Böhm *et al.*, *ibid*, p. 203.
10. S. Jadach *et al.*, *ibid*, p. 235.
11. R. Stuart, Phys. Lett. **B262**, 113 (1991).
12. A. Sirlin, Phys. Rev. Lett. **67**, 2127 (1991).
13. A. Leike, T. Riemann, and J. Rose, Phys. Lett. **B273**, 513 (1991).
14. See also D. Bardin *et al.*, Phys. Lett. **B206**, 539 (1988).
15. The LEP Collaborations: ALEPH, DELPHI, L3, OPAL, the LEP Electroweak Working Group, CERN-PH-EP/2013-022, **arXiv:1302.3415 [hep-ex]**, Phys.Rept. **532** (2013) 119-244.
16. W. Beenakker, F.A. Berends, and S.C. van der Marck, Nucl. Phys. **B349**, 323 (1991).
17. G. Montagna *et al.*, Nucl. Phys. **B401**, 3 (1993); Comp. Phys. Comm. **76**, 328 (1993); Comp. Phys. Comm. **93**,

# Gauge & Higgs Boson Particle Listings

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- 120 (1996);  
G. Montagna *et al.*, Comp. Phys. Comm. **117**, 278 (1999).
18. R. Assmann *et al.*, (Working Group on LEP Energy), Eur. Phys. J. **C6**, 187 (1999).
19. R. Assmann *et al.*, (Working Group on LEP Energy), Z. Phys. **C66**, 567 (1995).
20. L. Arnaudon *et al.*, (Working Group on LEP Energy and LEP Collabs.), Phys. Lett. **B307**, 187 (1993).
21. L. Arnaudon *et al.*, (Working Group on LEP Energy), CERN-PPE/92-125 (1992).
22. L. Arnaudon *et al.*, Phys. Lett. **B284**, 431 (1992).
23. R. Bailey *et al.*, 'LEP Energy Calibration' CERN-SL-90-95-AP, *Proceedings of the "2<sup>nd</sup> European Particle Accelerator Conference*, Nice, France, 12–16 June 1990, pp. 1765-1767.
24. The LEP Collabs.: ALEPH, DELPHI, L3, OPAL, the LEP Electroweak Working Group, and the SLD Heavy Flavour Group: Phys. Reports **427**, 257 (2006).
25. S. Jadach *et al.*, BHLUMI 4.04, Comp. Phys. Comm. **102**, 229 (1997);  
S. Jadach and O. Nicrosini, Event generators for Bhabha scattering, in Physics at LEP2, CERN-96-01 Vol. 2, February 1996.
26. B.F.L. Ward *et al.*, Phys. Lett. **B450**, 262 (1999).
27. W. Beenakker and G. Passarino, Phys. Lett. **B425**, 199 (1998).
28. M. Martinez *et al.*, Z. Phys. **C49**, 645 (1991);  
M. Martinez and F. Teubert, Z. Phys. **C65**, 267 (1995), updated with results summarized in S. Jadach, B. Pietrzyk, and M. Skrzypek, Phys. Lett. **B456**, 77 (1999) and Reports of the working group on precision calculations for the Z resonance, CERN 95-03, ed. D. Bardin, W. Hollik, and G. Passarino, and references therein.
29. T. van Ritbergen and R. Stuart, Phys. Lett. **B437**, 201 (1998); Phys. Rev. Lett. **82**, 488 (1999).
30. S. Eidelman and F. Jegerlehner, Z. Phys. **C67**, 585 (1995);  
M. Steinhäuser, Phys. Lett. **B429**, 158 (1998).
31. Particle Data Group (D.E. Groom *et al.*), Eur. Phys. J. **C15**, 1 (2000).
32. The LEP Experiments: ALEPH, DELPHI, L3, and OPAL Nucl. Instrum. Methods **A378**, 101 (1996).

### Z MASS

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note "The Z boson" and ref. LEP-SLC 06). The fit is performed using the Z mass and width, the Z hadronic pole cross section, the ratios of hadronic to leptonic partial widths, and the Z pole forward-backward lepton asymmetries. This set is believed to be most free of correlations.

The Z-boson mass listed here corresponds to the mass parameter in a Breit-Wigner distribution with mass dependent width. The value is 34 MeV greater than the real part of the position of the pole (in the energy-squared plane) in the Z-boson propagator. Also the LEP experiments have generally assumed a fixed value of the  $\gamma - Z$  interferences term based on the standard model. Keeping this term as free parameter leads to a somewhat larger error on the fitted Z mass. See ACCIARRI 00q and ABBIENDI 04g for a detailed investigation of both these issues.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>91.1876 ± 0.0021 OUR FIT</b>				
91.1852 ± 0.0030	4.57M	<sup>1</sup> ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
91.1863 ± 0.0028	4.08M	<sup>2</sup> ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
91.1898 ± 0.0031	3.96M	<sup>3</sup> ACCIARRI	00c L3	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
91.1885 ± 0.0031	4.57M	<sup>4</sup> BARATE	00c ALEP	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

91.1872 ± 0.0033		<sup>5</sup> ABBIENDI	04G OPAL	$E_{\text{cm}}^{\text{ee}} = \text{LEP1} +$ $130-209$ GeV
91.272 ± 0.032 ± 0.033		<sup>6</sup> ACHARD	04c L3	$E_{\text{cm}}^{\text{ee}} = 183-209$ GeV
91.1875 ± 0.0039	3.97M	<sup>7</sup> ACCIARRI	00Q L3	$E_{\text{cm}}^{\text{ee}} = \text{LEP1} +$ $130-189$ GeV
91.151 ± 0.008		<sup>8</sup> MIYABAYASHI	95 TOPZ	$E_{\text{cm}}^{\text{ee}} = 57.8$ GeV
91.74 ± 0.28 ± 0.93	156	<sup>9</sup> ALITTI	92B UA2	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 630$ GeV
90.9 ± 0.3 ± 0.2	188	<sup>10</sup> ABE	89c CDF	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 1.8$ TeV
91.14 ± 0.12	480	<sup>11</sup> ABRAMS	89B MRK2	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 89-93$ GeV
93.1 ± 1.0 ± 3.0	24	<sup>12</sup> ALBAJAR	89 UA1	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 546,630$ GeV

- <sup>1</sup> ABBIENDI 01A error includes approximately 2.3 MeV due to statistics and 1.8 MeV due to LEP energy uncertainty.
- <sup>2</sup> The error includes 1.6 MeV due to LEP energy uncertainty.
- <sup>3</sup> The error includes 1.8 MeV due to LEP energy uncertainty.
- <sup>4</sup> BARATE 00c error includes approximately 2.4 MeV due to statistics, 0.2 MeV due to experimental systematics, and 1.7 MeV due to LEP energy uncertainty.
- <sup>5</sup> ABBIENDI 04G obtain this result using the S-matrix formalism for a combined fit to their cross section and asymmetry data at the Z peak and their data at 130–209 GeV. The authors have corrected the measurement for the 34 MeV shift with respect to the Breit-Wigner fits.
- <sup>6</sup> ACHARD 04c select  $e^+e^- \rightarrow Z\gamma$  events with hard initial-state radiation. Z decays to  $q\bar{q}$  and muon pairs are considered. The fit results obtained in the two samples are found consistent to each other and combined considering the uncertainty due to ISR modelling as fully correlated.
- <sup>7</sup> ACCIARRI 00Q interpret the s-dependence of the cross sections and lepton forward-backward asymmetries in the framework of the S-matrix formalism. They fit to their cross section and asymmetry data at high energies, using the results of S-matrix fits to Z-peak data (ACCIARRI 00c) as constraints. The 130–189 GeV data constrains the  $\gamma/Z$  interference term. The authors have corrected the measurement for the 34.1 MeV shift with respect to the Breit-Wigner fits. The error contains a contribution of  $\pm 2.3$  MeV due to the uncertainty on the  $\gamma Z$  interference.
- <sup>8</sup> MIYABAYASHI 95 combine their low energy total hadronic cross-section measurement with the ACTON 93D data and perform a fit using an S-matrix formalism. As expected, this result is below the mass values obtained with the standard Breit-Wigner parametrization.
- <sup>9</sup> Enters fit through  $W/Z$  mass ratio given in the W Particle Listings. The ALITTI 92B systematic error ( $\pm 0.93$ ) has two contributions: one ( $\pm 0.92$ ) cancels in  $m_W/m_Z$  and one ( $\pm 0.12$ ) is noncancelling. These were added in quadrature.
- <sup>10</sup> First error of ABE 89 is combination of statistical and systematic contributions; second is mass scale uncertainty.
- <sup>11</sup> ABRAMS 89B uncertainty includes 35 MeV due to the absolute energy measurement.
- <sup>12</sup> ALBAJAR 89 result is from a total sample of  $33 Z \rightarrow e^+e^-$  events.

### Z WIDTH

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note "The Z boson" and ref. LEP-SLC 06).

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.4952 ± 0.0023 OUR FIT</b>				
2.4948 ± 0.0041	4.57M	<sup>1</sup> ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
2.4876 ± 0.0041	4.08M	<sup>2</sup> ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
2.5024 ± 0.0042	3.96M	<sup>3</sup> ACCIARRI	00c L3	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
2.4951 ± 0.0043	4.57M	<sup>4</sup> BARATE	00c ALEP	$E_{\text{cm}}^{\text{ee}} = 88-94$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.4943 ± 0.0041		<sup>5</sup> ABBIENDI	04G OPAL	$E_{\text{cm}}^{\text{ee}} = \text{LEP1} +$ $130-209$ GeV
2.5025 ± 0.0041	3.97M	<sup>6</sup> ACCIARRI	00Q L3	$E_{\text{cm}}^{\text{ee}} = \text{LEP1} +$ $130-189$ GeV
2.50 ± 0.21 ± 0.06		<sup>7</sup> ABREU	96R DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV
3.8 ± 0.8 ± 1.0	188	<sup>8</sup> ABE	89c CDF	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 1.8$ TeV
2.42 $^{+0.45}_{-0.35}$	480	<sup>9</sup> ABRAMS	89B MRK2	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 89-93$ GeV
2.7 $^{+1.2}_{-1.0}$ ± 1.3	24	<sup>10</sup> ALBAJAR	89 UA1	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 546,630$ GeV
2.7 ± 2.0 ± 1.0	25	<sup>10</sup> ANSARI	87 UA2	$E_{\text{cm}}^{\text{e}\bar{\text{e}}} = 546,630$ GeV

- <sup>1</sup> ABBIENDI 01A error includes approximately 3.6 MeV due to statistics, 1 MeV due to event selection systematics, and 1.3 MeV due to LEP energy uncertainty.
- <sup>2</sup> The error includes 1.2 MeV due to LEP energy uncertainty.
- <sup>3</sup> The error includes 1.3 MeV due to LEP energy uncertainty.
- <sup>4</sup> BARATE 00c error includes approximately 3.8 MeV due to statistics, 0.9 MeV due to experimental systematics, and 1.3 MeV due to LEP energy uncertainty.
- <sup>5</sup> ABBIENDI 04G obtain this result using the S-matrix formalism for a combined fit to their cross section and asymmetry data at the Z peak and their data at 130–209 GeV. The authors have corrected the measurement for the 1 MeV shift with respect to the Breit-Wigner fits.
- <sup>6</sup> ACCIARRI 00Q interpret the s-dependence of the cross sections and lepton forward-backward asymmetries in the framework of the S-matrix formalism. They fit to their cross section and asymmetry data at high energies, using the results of S-matrix fits to Z-peak data (ACCIARRI 00c) as constraints. The 130–189 GeV data constrains the  $\gamma/Z$  interference term. The authors have corrected the measurement for the 0.9 MeV shift with respect to the Breit-Wigner fits.
- <sup>7</sup> ABREU 96R obtain this value from a study of the interference between initial and final state radiation in the process  $e^+e^- \rightarrow Z \rightarrow \mu^+\mu^-$ .
- <sup>8</sup> ABRAMS 89B uncertainty includes 50 MeV due to the miniSAM background subtraction error.
- <sup>9</sup> ALBAJAR 89 result is from a total sample of  $33 Z \rightarrow e^+e^-$  events.

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<sup>10</sup> Quoted values of ANSARI 87 are from direct fit. Ratio of Z and W production gives either  $\Gamma(Z) < (1.09 \pm 0.07) \times \Gamma(W)$ , CL = 90% or  $\Gamma(Z) = (0.82^{+0.19}_{-0.14} \pm 0.06) \times \Gamma(W)$ . Assuming Standard-Model value  $\Gamma(W) = 2.65$  GeV then gives  $\Gamma(Z) < 2.89 \pm 0.19$  or  $= 2.17^{+0.50}_{-0.37} \pm 0.16$ .

### Z DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $e^+ e^-$	( 3.363 $\pm$ 0.004 ) %	
$\Gamma_2$ $\mu^+ \mu^-$	( 3.366 $\pm$ 0.007 ) %	
$\Gamma_3$ $\tau^+ \tau^-$	( 3.370 $\pm$ 0.008 ) %	
$\Gamma_4$ $\ell^+ \ell^-$	[a] ( 3.3658 $\pm$ 0.0023 ) %	
$\Gamma_5$ $\ell^+ \ell^- \ell^+ \ell^-$	[b] ( 3.30 $\pm$ 0.31 ) $\times 10^{-6}$	S=1.1
$\Gamma_6$ invisible	(20.00 $\pm$ 0.06 ) %	
$\Gamma_7$ hadrons	(69.91 $\pm$ 0.06 ) %	
$\Gamma_8$ $(u\bar{u} + c\bar{c})/2$	(11.6 $\pm$ 0.6 ) %	
$\Gamma_9$ $(d\bar{d} + s\bar{s} + b\bar{b})/3$	(15.6 $\pm$ 0.4 ) %	
$\Gamma_{10}$ $c\bar{c}$	(12.03 $\pm$ 0.21 ) %	
$\Gamma_{11}$ $b\bar{b}$	(15.12 $\pm$ 0.05 ) %	
$\Gamma_{12}$ $b\bar{b}b\bar{b}$	( 3.6 $\pm$ 1.3 ) $\times 10^{-4}$	
$\Gamma_{13}$ $ggg$	< 1.1 %	CL=95%
$\Gamma_{14}$ $\pi^0 \gamma$	< 2.01 $\times 10^{-5}$	CL=95%
$\Gamma_{15}$ $\eta \gamma$	< 5.1 $\times 10^{-5}$	CL=95%
$\Gamma_{16}$ $\omega \gamma$	< 6.5 $\times 10^{-4}$	CL=95%
$\Gamma_{17}$ $\eta'(958) \gamma$	< 4.2 $\times 10^{-5}$	CL=95%
$\Gamma_{18}$ $\gamma \gamma$	< 1.46 $\times 10^{-5}$	CL=95%
$\Gamma_{19}$ $\pi^0 \pi^0$	< 1.52 $\times 10^{-5}$	CL=95%
$\Gamma_{20}$ $\gamma \gamma \gamma$	< 1.0 $\times 10^{-5}$	CL=95%
$\Gamma_{21}$ $\pi^\pm W^\mp$	[c] < 7 $\times 10^{-5}$	CL=95%
$\Gamma_{22}$ $\rho^\pm W^\mp$	[c] < 8.3 $\times 10^{-5}$	CL=95%
$\Gamma_{23}$ $J/\psi(1S) X$	( 3.51 $^{+0.23}_{-0.25}$ ) $\times 10^{-3}$	S=1.1
$\Gamma_{24}$ $J/\psi(1S) \gamma$	< 2.6 $\times 10^{-6}$	CL=95%
$\Gamma_{25}$ $\psi(2S) X$	( 1.60 $\pm$ 0.29 ) $\times 10^{-3}$	
$\Gamma_{26}$ $\chi_{c1}(1P) X$	( 2.9 $\pm$ 0.7 ) $\times 10^{-3}$	
$\Gamma_{27}$ $\chi_{c2}(1P) X$	< 3.2 $\times 10^{-3}$	CL=90%
$\Gamma_{28}$ $\Upsilon(1S) X + \Upsilon(2S) X + \Upsilon(3S) X$	( 1.0 $\pm$ 0.5 ) $\times 10^{-4}$	
$\Gamma_{29}$ $\Upsilon(1S) X$	< 3.4 $\times 10^{-6}$	CL=95%
$\Gamma_{30}$ $\Upsilon(2S) X$	< 6.5 $\times 10^{-6}$	CL=95%
$\Gamma_{31}$ $\Upsilon(3S) X$	< 5.4 $\times 10^{-6}$	CL=95%
$\Gamma_{32}$ $(D^0/\bar{D}^0) X$	(20.7 $\pm$ 2.0 ) %	
$\Gamma_{33}$ $D^\pm X$	(12.2 $\pm$ 1.7 ) %	
$\Gamma_{34}$ $D^*(2010)^\pm X$	[c] (11.4 $\pm$ 1.3 ) %	
$\Gamma_{35}$ $D_{s1}(2536)^\pm X$	( 3.6 $\pm$ 0.8 ) $\times 10^{-3}$	
$\Gamma_{36}$ $D_{sJ}(2573)^\pm X$	( 5.8 $\pm$ 2.2 ) $\times 10^{-3}$	
$\Gamma_{37}$ $D_s^*(2629)^\pm X$	searched for	
$\Gamma_{38}$ $B X$		
$\Gamma_{39}$ $B^* X$		
$\Gamma_{40}$ $B^+ X$	[d] ( 6.08 $\pm$ 0.13 ) %	
$\Gamma_{41}$ $B_s^0 X$	[d] ( 1.59 $\pm$ 0.13 ) %	
$\Gamma_{42}$ $B_c^\pm X$	searched for	
$\Gamma_{43}$ $\Lambda_c^+ X$	( 1.54 $\pm$ 0.33 ) %	
$\Gamma_{44}$ $\Xi_c^0 X$	seen	
$\Gamma_{45}$ $\Xi_b X$	seen	
$\Gamma_{46}$ $b$ -baryon X	[d] ( 1.38 $\pm$ 0.22 ) %	
$\Gamma_{47}$ anomalous $\gamma$ + hadrons	[e] < 3.2 $\times 10^{-3}$	CL=95%
$\Gamma_{48}$ $e^+ e^- \gamma$	[e] < 5.2 $\times 10^{-4}$	CL=95%
$\Gamma_{49}$ $\mu^+ \mu^- \gamma$	[e] < 5.6 $\times 10^{-4}$	CL=95%
$\Gamma_{50}$ $\tau^+ \tau^- \gamma$	[e] < 7.3 $\times 10^{-4}$	CL=95%
$\Gamma_{51}$ $\ell^+ \ell^- \gamma \gamma$	[f] < 6.8 $\times 10^{-6}$	CL=95%
$\Gamma_{52}$ $q\bar{q} \gamma \gamma$	[f] < 5.5 $\times 10^{-6}$	CL=95%
$\Gamma_{53}$ $\nu\bar{\nu} \gamma \gamma$	[f] < 3.1 $\times 10^{-6}$	CL=95%
$\Gamma_{54}$ $e^\pm \mu^\mp$	LF [c] < 7.5 $\times 10^{-7}$	CL=95%
$\Gamma_{55}$ $e^\pm \tau^\mp$	LF [c] < 9.8 $\times 10^{-6}$	CL=95%
$\Gamma_{56}$ $\mu^\pm \tau^\mp$	LF [c] < 1.2 $\times 10^{-5}$	CL=95%
$\Gamma_{57}$ $p e$	L,B < 1.8 $\times 10^{-6}$	CL=95%
$\Gamma_{58}$ $p \mu$	L,B < 1.8 $\times 10^{-6}$	CL=95%

[a]  $\ell$  indicates each type of lepton ( $e$ ,  $\mu$ , and  $\tau$ ), not sum over them.

[b] Here  $\ell$  indicates  $e$  or  $\mu$ .

[c] The value is for the sum of the charge states or particle/antiparticle states indicated.

[d] This value is updated using the product of (i) the  $Z \rightarrow b\bar{b}$  fraction from this listing and (ii) the  $b$ -hadron fraction in an unbiased sample of weakly decaying  $b$ -hadrons produced in Z-decays provided by the Heavy Flavor Averaging Group (HFAG, [http://www.slac.stanford.edu/xorg/hfag/osc/PDG\\_2009/#FRACZ](http://www.slac.stanford.edu/xorg/hfag/osc/PDG_2009/#FRACZ)).

[e] See the Particle Listings below for the  $\gamma$  energy range used in this measurement.

[f] For  $m_{\gamma\gamma} = (60 \pm 5)$  GeV.

### Z PARTIAL WIDTHS

#### $\Gamma(e^+ e^-)$

$\Gamma_1$

For the LEP experiments, this parameter is not directly used in the overall fit but is derived using the fit results; see the note "The Z boson" and ref. LEP-SLC 06.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>83.91 <math>\pm</math> 0.12 OUR FIT</b>				
83.66 $\pm$ 0.20	137.0K	ABBIENDI	01A OPAL	$E_{cm}^{ee} = 88-94$ GeV
83.54 $\pm$ 0.27	117.8k	ABREU	00F DLPH	$E_{cm}^{ee} = 88-94$ GeV
84.16 $\pm$ 0.22	124.4k	ACCIARRI	00C L3	$E_{cm}^{ee} = 88-94$ GeV
83.88 $\pm$ 0.19		BARATE	00C ALEP	$E_{cm}^{ee} = 88-94$ GeV
82.89 $\pm$ 1.20 $\pm$ 0.89		<sup>1</sup> ABE	95J SLD	$E_{cm}^{ee} = 91.31$ GeV

<sup>1</sup> ABE 95J obtain this measurement from Bhabha events in a restricted fiducial region to improve systematics. They use the values 91.187 and 2.489 GeV for the Z mass and total decay width to extract this partial width.

#### $\Gamma(\mu^+ \mu^-)$

$\Gamma_2$

This parameter is not directly used in the overall fit but is derived using the fit results; see the note "The Z boson" and ref. LEP-SLC 06.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>83.99 <math>\pm</math> 0.18 OUR FIT</b>				
84.03 $\pm$ 0.30	182.8K	ABBIENDI	01A OPAL	$E_{cm}^{ee} = 88-94$ GeV
84.48 $\pm$ 0.40	157.6k	ABREU	00F DLPH	$E_{cm}^{ee} = 88-94$ GeV
83.95 $\pm$ 0.44	113.4k	ACCIARRI	00C L3	$E_{cm}^{ee} = 88-94$ GeV
84.02 $\pm$ 0.28		BARATE	00C ALEP	$E_{cm}^{ee} = 88-94$ GeV

#### $\Gamma(\tau^+ \tau^-)$

$\Gamma_3$

This parameter is not directly used in the overall fit but is derived using the fit results; see the note "The Z boson" and ref. LEP-SLC 06.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>84.08 <math>\pm</math> 0.22 OUR FIT</b>				
83.94 $\pm$ 0.41	151.5K	ABBIENDI	01A OPAL	$E_{cm}^{ee} = 88-94$ GeV
83.71 $\pm$ 0.58	104.0k	ABREU	00F DLPH	$E_{cm}^{ee} = 88-94$ GeV
84.23 $\pm$ 0.58	103.0k	ACCIARRI	00C L3	$E_{cm}^{ee} = 88-94$ GeV
84.38 $\pm$ 0.31		BARATE	00C ALEP	$E_{cm}^{ee} = 88-94$ GeV

#### $\Gamma(\ell^+ \ell^-)$

$\Gamma_4$

In our fit  $\Gamma(\ell^+ \ell^-)$  is defined as the partial Z width for the decay into a pair of massless charged leptons. This parameter is not directly used in the 5-parameter fit assuming lepton universality but is derived using the fit results. See the note "The Z boson" and ref. LEP-SLC 06.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>83.984 <math>\pm</math> 0.086 OUR FIT</b>				
83.82 $\pm$ 0.15	471.3K	ABBIENDI	01A OPAL	$E_{cm}^{ee} = 88-94$ GeV
83.85 $\pm$ 0.17	379.4k	ABREU	00F DLPH	$E_{cm}^{ee} = 88-94$ GeV
84.14 $\pm$ 0.17	340.8k	ACCIARRI	00C L3	$E_{cm}^{ee} = 88-94$ GeV
84.02 $\pm$ 0.15	500k	BARATE	00C ALEP	$E_{cm}^{ee} = 88-94$ GeV

#### $\Gamma(\text{invisible})$

$\Gamma_6$

We use only direct measurements of the invisible partial width using the single photon channel to obtain the average value quoted below. OUR FIT value is obtained as a difference between the total and the observed partial widths assuming lepton universality.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>499.0 <math>\pm</math> 1.5 OUR FIT</b>				
<b>503 <math>\pm</math> 16 OUR AVERAGE</b>	Error	includes scale factor of 1.2.		
498 $\pm$ 12 $\pm$ 12	1791	ACCIARRI	98G L3	$E_{cm}^{ee} = 88-94$ GeV
539 $\pm$ 26 $\pm$ 17	410	AKERS	95C OPAL	$E_{cm}^{ee} = 88-94$ GeV
450 $\pm$ 34 $\pm$ 34	258	BUSKULIC	93L ALEP	$E_{cm}^{ee} = 88-94$ GeV
540 $\pm$ 80 $\pm$ 40	52	ADEVA	92 L3	$E_{cm}^{ee} = 88-94$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

498.1 $\pm$ 2.6	<sup>1</sup> ABBIENDI	01A OPAL	$E_{cm}^{ee} = 88-94$ GeV
498.1 $\pm$ 3.2	<sup>1</sup> ABREU	00F DLPH	$E_{cm}^{ee} = 88-94$ GeV
499.1 $\pm$ 2.9	<sup>1</sup> ACCIARRI	00C L3	$E_{cm}^{ee} = 88-94$ GeV
499.1 $\pm$ 2.5	<sup>1</sup> BARATE	00C ALEP	$E_{cm}^{ee} = 88-94$ GeV

<sup>1</sup> This is an indirect determination of  $\Gamma(\text{invisible})$  from a fit to the visible Z decay modes.

# Gauge & Higgs Boson Particle Listings

## Z

**$\Gamma(\text{hadrons})$**   **$\Gamma_7$**   
This parameter is not directly used in the 5-parameter fit assuming lepton universality, but is derived using the fit results. See the note “The Z boson” and ref. LEP-SLC 06.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1744.4±2.0 OUR FIT</b>				
1745.4±3.5	4.10M	ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
1738.1±4.0	3.70M	ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
1751.1±3.8	3.54M	ACCIARRI	00C L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
1744.0±3.4	4.07M	BARATE	00C ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

### Z BRANCHING RATIOS

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note “The Z boson” and ref. LEP-SLC 06).

<b><math>\Gamma(\text{hadrons})/\Gamma(e^+e^-)</math></b>	<b><math>\Gamma_7/\Gamma_1</math></b>			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>20.804±0.050 OUR FIT</b>				
20.902±0.084	137.0K	<sup>1</sup> ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.88 ±0.12	117.8k	ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.816±0.089	124.4k	ACCIARRI	00C L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.677±0.075		<sup>2</sup> BARATE	00C ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
27.0 $\begin{smallmatrix} +11.7 \\ -8.8 \end{smallmatrix}$	12	<sup>3</sup> ABRAMS	89D MRK2	$E_{\text{cm}}^{\text{ee}} = 89\text{--}93$ GeV

- <sup>1</sup> ABBIENDI 01A error includes approximately 0.067 due to statistics, 0.040 due to event selection systematics, 0.027 due to the theoretical uncertainty in  $t$ -channel prediction, and 0.014 due to LEP energy uncertainty.
- <sup>2</sup> BARATE 00C error includes approximately 0.062 due to statistics, 0.033 due to experimental systematics, and 0.026 due to the theoretical uncertainty in  $t$ -channel prediction.
- <sup>3</sup> ABRAMS 89D have included both statistical and systematic uncertainties in their quoted errors.

**$\Gamma(\text{hadrons})/\Gamma(\mu^+\mu^-)$**   **$\Gamma_7/\Gamma_2$**   
OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note “The Z boson” and ref. LEP-SLC 06).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>20.785±0.033 OUR FIT</b>				
20.811±0.058	182.8K	<sup>1</sup> ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.65 ±0.08	157.6k	ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.861±0.097	113.4k	ACCIARRI	00C L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.799±0.056		<sup>2</sup> BARATE	00C ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
18.9 $\begin{smallmatrix} +7.1 \\ -5.3 \end{smallmatrix}$	13	<sup>3</sup> ABRAMS	89D MRK2	$E_{\text{cm}}^{\text{ee}} = 89\text{--}93$ GeV

- <sup>1</sup> ABBIENDI 01A error includes approximately 0.050 due to statistics and 0.027 due to event selection systematics.
- <sup>2</sup> BARATE 00C error includes approximately 0.053 due to statistics and 0.021 due to experimental systematics.
- <sup>3</sup> ABRAMS 89D have included both statistical and systematic uncertainties in their quoted errors.

**$\Gamma(\text{hadrons})/\Gamma(\tau^+\tau^-)$**   **$\Gamma_7/\Gamma_3$**   
OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note “The Z boson” and ref. LEP-SLC 06).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>20.764±0.045 OUR FIT</b>				
20.832±0.091	151.5K	<sup>1</sup> ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.84 ±0.13	104.0k	ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.792±0.133	103.0k	ACCIARRI	00C L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.707±0.062		<sup>2</sup> BARATE	00C ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
15.2 $\begin{smallmatrix} +4.8 \\ -3.9 \end{smallmatrix}$	21	<sup>3</sup> ABRAMS	89D MRK2	$E_{\text{cm}}^{\text{ee}} = 89\text{--}93$ GeV

- <sup>1</sup> ABBIENDI 01A error includes approximately 0.055 due to statistics and 0.071 due to event selection systematics.
- <sup>2</sup> BARATE 00C error includes approximately 0.054 due to statistics and 0.033 due to experimental systematics.
- <sup>3</sup> ABRAMS 89D have included both statistical and systematic uncertainties in their quoted errors.

**$\Gamma(\text{hadrons})/\Gamma(\ell^+\ell^-)$**   **$\Gamma_7/\Gamma_4$**   
 $\ell$  indicates each type of lepton ( $e$ ,  $\mu$ , and  $\tau$ ), not sum over them.

Our fit result is obtained requiring lepton universality.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>20.767±0.025 OUR FIT</b>				
20.823±0.044	471.3K	<sup>1</sup> ABBIENDI	01A OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.730±0.060	379.4k	ABREU	00F DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.810±0.060	340.8k	ACCIARRI	00C L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
20.725±0.039	500k	<sup>2</sup> BARATE	00C ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
18.9 $\begin{smallmatrix} +3.6 \\ -3.2 \end{smallmatrix}$	46	ABRAMS	89B MRK2	$E_{\text{cm}}^{\text{ee}} = 89\text{--}93$ GeV

- <sup>1</sup> ABBIENDI 01A error includes approximately 0.034 due to statistics and 0.027 due to event selection systematics.
- <sup>2</sup> BARATE 00C error includes approximately 0.033 due to statistics, 0.020 due to experimental systematics, and 0.005 due to the theoretical uncertainty in  $t$ -channel prediction.

**$\Gamma(\text{hadrons})/\Gamma_{\text{total}}$**   **$\Gamma_7/\Gamma$**   
This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE (%)	DOCUMENT ID
<b>69.911±0.056 OUR FIT</b>	

**$\Gamma(e^+e^-)/\Gamma_{\text{total}}$**   **$\Gamma_1/\Gamma$**   
This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE (%)	DOCUMENT ID
<b>3.3632±0.0042 OUR FIT</b>	

**$\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$**   **$\Gamma_2/\Gamma$**   
This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE (%)	DOCUMENT ID
<b>3.3662±0.0066 OUR FIT</b>	

**$\Gamma(\mu^+\mu^-)/\Gamma(e^+e^-)$**   **$\Gamma_2/\Gamma_1$**   
This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE	DOCUMENT ID
<b>1.0009±0.0028 OUR FIT</b>	

**$\Gamma(\tau^+\tau^-)/\Gamma_{\text{total}}$**   **$\Gamma_3/\Gamma$**   
This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE (%)	DOCUMENT ID
<b>3.3696±0.0083 OUR FIT</b>	

**$\Gamma(\tau^+\tau^-)/\Gamma(e^+e^-)$**   **$\Gamma_3/\Gamma_1$**   
This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE	DOCUMENT ID
<b>1.0019±0.0032 OUR FIT</b>	

**$\Gamma(\ell^+\ell^-)/\Gamma_{\text{total}}$**   **$\Gamma_4/\Gamma$**   
 $\ell$  indicates each type of lepton ( $e$ ,  $\mu$ , and  $\tau$ ), not sum over them.

Our fit result assumes lepton universality.

This parameter is not directly used in the overall fit but is derived using the fit results; see the note “The Z boson” and ref. LEP-SLC 06.

VALUE (%)	DOCUMENT ID
<b>3.3658±0.0023 OUR FIT</b>	

**$\Gamma(\ell^+\ell^-\ell^+\ell^-)/\Gamma_{\text{total}}$**   **$\Gamma_5/\Gamma$**   
Here  $\ell$  indicates either  $e$  or  $\mu$ .

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.30±0.31 OUR AVERAGE</b>				Error includes scale factor of 1.1.
3.20±0.25±0.13	172	AAD	14N ATLS	$E_{\text{cm}}^{\text{pp}} = 7, 8$ TeV
4.2 $\begin{smallmatrix} +0.9 \\ -0.8 \end{smallmatrix}$ ±0.2	28	CHATRCHYAN12B	CMS	$E_{\text{cm}}^{\text{pp}} = 7$ TeV

**$\Gamma(\text{invisible})/\Gamma_{\text{total}}$**   **$\Gamma_6/\Gamma$**   
See the data, the note, and the fit result for the partial width,  $\Gamma_6$ , above.

VALUE (%)	DOCUMENT ID
<b>20.000±0.055 OUR FIT</b>	

**$\Gamma((u\bar{u}+c\bar{c})/2)/\Gamma(\text{hadrons})$**   **$\Gamma_8/\Gamma_7$**   
This quantity is the branching ratio of  $Z \rightarrow$  “up-type” quarks to  $Z \rightarrow$  hadrons. Except ACKERSTAFF 97T the values of  $Z \rightarrow$  “up-type” and  $Z \rightarrow$  “down-type” branchings are extracted from measurements of  $\Gamma(\text{hadrons})$ , and  $\Gamma(Z \rightarrow \gamma + \text{jets})$  where  $\gamma$  is a high-energy ( $>5$  or  $7$  GeV) isolated photon. As the experiments use different procedures and slightly different values of  $M_Z$ ,  $\Gamma(\text{hadrons})$  and  $\alpha_s$  in their extraction procedures, our average has to be taken with caution.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.166±0.009 OUR AVERAGE</b>			
0.172 $\begin{smallmatrix} +0.011 \\ -0.010 \end{smallmatrix}$	<sup>1</sup> ABBIENDI	04E OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV
0.160±0.019±0.019	<sup>2</sup> ACKERSTAFF 97T	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.137±0.038 $\begin{smallmatrix} +0.025 \\ -0.054 \end{smallmatrix}$	<sup>3</sup> ABREU	95x DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.137±0.033	<sup>4</sup> ADRIANI	93 L3	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV

- <sup>1</sup> ABBIENDI 04E select photons with energy  $> 7$  GeV and use  $\Gamma(\text{hadrons}) = 1744.4 \pm 2.0$  MeV and  $\alpha_s = 0.1172 \pm 0.002$  to obtain  $\Gamma_u = 300^{+19}_{-18}$  MeV.
- <sup>2</sup> ACKERSTAFF 97T measure  $\Gamma_{u\bar{u}}/(\Gamma_{d\bar{d}}+\Gamma_{u\bar{u}}+\Gamma_{s\bar{s}}) = 0.258 \pm 0.031 \pm 0.032$ . To obtain this branching ratio authors use  $R_c+R_b = 0.380 \pm 0.010$ . This measurement is fully negatively correlated with the measurement of  $\Gamma_{d\bar{d},s\bar{s}}/(\Gamma_{d\bar{d}}+\Gamma_{u\bar{u}}+\Gamma_{s\bar{s}})$  given in the next data block.
- <sup>3</sup> ABREU 95x use  $M_Z = 91.187 \pm 0.009$  GeV,  $\Gamma(\text{hadrons}) = 1725 \pm 12$  MeV and  $\alpha_s = 0.123 \pm 0.005$ . To obtain this branching ratio we divide their value of  $C_{2/3} = 0.91^{+0.25}_{-0.36}$  by their value of  $(3C_{1/3} + 2C_{2/3}) = 6.66 \pm 0.05$ .
- <sup>4</sup> ADRIANI 93 use  $M_Z = 91.181 \pm 0.022$  GeV,  $\Gamma(\text{hadrons}) = 1742 \pm 19$  MeV and  $\alpha_s = 0.125 \pm 0.009$ . To obtain this branching ratio we divide their value of  $C_{2/3} = 0.92 \pm 0.22$  by their value of  $(3C_{1/3} + 2C_{2/3}) = 6.720 \pm 0.076$ .



$\Gamma((d\bar{d} + s\bar{s} + b\bar{b})/3)/\Gamma(\text{hadrons})$  $\Gamma_9/\Gamma_7$ 

This quantity is the branching ratio of  $Z \rightarrow$  “down-type” quarks to  $Z \rightarrow$  hadrons. Except ACKERSTAFF 97T the values of  $Z \rightarrow$  “up-type” and  $Z \rightarrow$  “down-type” branchings are extracted from measurements of  $\Gamma(\text{hadrons})$ , and  $\Gamma(Z \rightarrow \gamma + \text{jets})$  where  $\gamma$  is a high-energy ( $>5$  or  $7$  GeV) isolated photon. As the experiments use different procedures and slightly different values of  $M_Z$ ,  $\Gamma(\text{hadrons})$  and  $\alpha_s$  in their extraction procedures, our average has to be taken with caution.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.223 ± 0.006 OUR AVERAGE</b>			
0.218 ± 0.007	<sup>1</sup> ABBIENDI	04E	OPAL $E_{\text{cm}}^{\text{ee}} = 91.2$ GeV
0.230 ± 0.010 ± 0.010	<sup>2</sup> ACKERSTAFF	97T	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.243 $\pm 0.036$ $-0.026$	<sup>3</sup> ABREU	95x	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.243 ± 0.022	<sup>4</sup> ADRIANI	93	L3 $E_{\text{cm}}^{\text{ee}} = 91.2$ GeV

<sup>1</sup> ABBIENDI 04E select photons with energy  $> 7$  GeV and use  $\Gamma(\text{hadrons}) = 1744.4 \pm 2.0$  MeV and  $\alpha_s = 0.1172 \pm 0.002$  to obtain  $\Gamma_d = 381 \pm 12$  MeV.

<sup>2</sup> ACKERSTAFF 97T measure  $\Gamma_{d\bar{d}s\bar{s}}/(\Gamma_{d\bar{d}} + \Gamma_{u\bar{u}} + \Gamma_{s\bar{s}}) = 0.371 \pm 0.016 \pm 0.016$ . To obtain this branching ratio authors use  $R_C + R_b = 0.380 \pm 0.010$ . This measurement is fully negatively correlated with the measurement of  $\Gamma_{u\bar{u}}/(\Gamma_{d\bar{d}} + \Gamma_{u\bar{u}} + \Gamma_{s\bar{s}})$  presented in the previous data block.

<sup>3</sup> ABREU 95x use  $M_Z = 91.187 \pm 0.009$  GeV,  $\Gamma(\text{hadrons}) = 1725 \pm 12$  MeV and  $\alpha_s = 0.123 \pm 0.005$ . To obtain this branching ratio we divide their value of  $C_{1/3} = 1.62 \pm 0.24$  by their value of  $(3C_{1/3} + 2C_{2/3}) = 6.66 \pm 0.05$ .

<sup>4</sup> ADRIANI 93 use  $M_Z = 91.181 \pm 0.022$  GeV,  $\Gamma(\text{hadrons}) = 1742 \pm 19$  MeV and  $\alpha_s = 0.125 \pm 0.009$ . To obtain this branching ratio we divide their value of  $C_{1/3} = 1.63 \pm 0.15$  by their value of  $(3C_{1/3} + 2C_{2/3}) = 6.720 \pm 0.076$ .

 $R_C = \Gamma(c\bar{c})/\Gamma(\text{hadrons})$  $\Gamma_{10}/\Gamma_7$ 

OUR FIT is obtained by a simultaneous fit to several  $c$ - and  $b$ -quark measurements as explained in the note “The Z boson” and ref. LEP-SLC 06.

The Standard Model predicts  $R_C = 0.1723$  for  $m_t = 174.3$  GeV and  $M_H = 150$  GeV.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.1721 ± 0.0030 OUR FIT</b>			
0.1744 ± 0.0031 ± 0.0021	<sup>1</sup> ABE	05F	SLD $E_{\text{cm}}^{\text{ee}} = 91.28$ GeV
0.1665 ± 0.0051 ± 0.0081	<sup>2</sup> ABREU	00	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.1698 ± 0.0069	<sup>3</sup> BARATE	00B	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.180 ± 0.011 ± 0.013	<sup>4</sup> ACKERSTAFF	98E	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.167 ± 0.011 ± 0.012	<sup>5</sup> ALEXANDER	96R	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1623 ± 0.0085 ± 0.0209	<sup>6</sup> ABREU	95D	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ABE 05F use hadronic Z decays collected during 1996–98 to obtain an enriched sample of  $c\bar{c}$  events using a double tag method. The single  $c$ -tag is obtained with a neural network trained to perform flavor discrimination using as input several signatures (corrected secondary vertex mass, vertex decay length, multiplicity and total momentum of the hemisphere). A multitag approach is used, defining 4 regions of the output value of the neural network and  $R_C$  is extracted from a simultaneous fit to the count rates of the 4 different tags. The quoted systematic error includes an uncertainty of  $\pm 0.0006$  due to the uncertainty on  $R_b$ .

<sup>2</sup> ABREU 00 obtain this result properly combining the measurement from the  $D^{*+}$  production rate ( $R_C = 0.1610 \pm 0.0104 \pm 0.0077 \pm 0.0043$  (BR)) with that from the overall charm counting ( $R_C = 0.1692 \pm 0.0047 \pm 0.0063 \pm 0.0074$  (BR)) in  $c\bar{c}$  events. The systematic error includes an uncertainty of  $\pm 0.0054$  due to the uncertainty on the charmed hadron branching fractions.

<sup>3</sup> BARATE 00B use exclusive decay modes to independently determine the quantities  $R_C \times f(c \rightarrow X)$ ,  $X = D^0$ ,  $D^+$ ,  $D_s^+$ , and  $\Lambda_c$ . Estimating  $R_C \times f(c \rightarrow \Xi_c^-/\Omega_c^-) = 0.0034$ , they simply sum over all the charm decays to obtain  $R_C = 0.1738 \pm 0.0047 \pm 0.0088 \pm 0.0075$  (BR). This is combined with all previous ALEPH measurements (BARATE 98T and BUSKULIC 94G,  $R_C = 0.1681 \pm 0.0054 \pm 0.0062$ ) to obtain the quoted value.

<sup>4</sup> ACKERSTAFF 98E use an inclusive/exclusive double tag. In one jet  $D^{*\pm}$  mesons are exclusively reconstructed in several decay channels and in the opposite jet a slow pion (opposite charge inclusive  $D^{*\pm}$ ) tag is used. The  $b$  content of this sample is measured by the simultaneous detection of a lepton in one jet and an inclusively reconstructed  $D^{*\pm}$  meson in the opposite jet. The systematic error includes an uncertainty of  $\pm 0.006$  due to the external branching ratios.

<sup>5</sup> ALEXANDER 96R obtain this value via direct charm counting, summing the partial contributions from  $D^0$ ,  $D^+$ ,  $D_s^+$ , and  $\Lambda_c^+$ , and assuming that strange-charmed baryons account for the 15% of the  $\Lambda_c^+$  production. An uncertainty of  $\pm 0.005$  due to the uncertainties in the charm hadron branching ratios is included in the overall systematics.

<sup>6</sup> ABREU 95D perform a maximum likelihood fit to the combined  $p$  and  $p_T$  distributions of single and dilepton samples. The second error includes an uncertainty of  $\pm 0.0124$  due to models and branching ratios.

 $R_b = \Gamma(b\bar{b})/\Gamma(\text{hadrons})$  $\Gamma_{11}/\Gamma_7$ 

OUR FIT is obtained by a simultaneous fit to several  $c$ - and  $b$ -quark measurements as explained in the note “The Z boson” and ref. LEP-SLC 06.

The Standard Model predicts  $R_b = 0.21581$  for  $m_t = 174.3$  GeV and  $M_H = 150$  GeV.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.21629 ± 0.00066 OUR FIT</b>			
0.21594 ± 0.00094 ± 0.00075	<sup>1</sup> ABE	05F	SLD $E_{\text{cm}}^{\text{ee}} = 91.28$ GeV
0.2174 ± 0.0015 ± 0.0028	<sup>2</sup> ACCIARRI	00	L3 $E_{\text{cm}}^{\text{ee}} = 89\text{--}93$ GeV
0.2178 ± 0.0011 ± 0.0013	<sup>3</sup> ABBIENDI	99B	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.21634 ± 0.00067 ± 0.00060	<sup>4</sup> ABREU	99B	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.2159 ± 0.0009 ± 0.0011	<sup>5</sup> BARATE	97F	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.2145 ± 0.0089 ± 0.0067	<sup>6</sup> ABREU	95D	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.219 ± 0.006 ± 0.005	<sup>7</sup> BUSKULIC	94G	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.251 ± 0.049 ± 0.030	<sup>8</sup> JACOBSEN	91	MRK2 $E_{\text{cm}}^{\text{ee}} = 91$ GeV

<sup>1</sup> ABE 05F use hadronic Z decays collected during 1996–98 to obtain an enriched sample of  $b\bar{b}$  events using a double tag method. The single  $b$ -tag is obtained with a neural network trained to perform flavor discrimination using as input several signatures (corrected secondary vertex mass, vertex decay length, multiplicity and total momentum of the hemisphere; the key tag is obtained requiring the secondary vertex corrected mass to be above the  $D$ -meson mass). ABE 05F obtain  $R_b = 0.21604 \pm 0.00098 \pm 0.00074$  where the systematic error includes an uncertainty of  $\pm 0.00012$  due to the uncertainty on  $R_C$ . The value reported here is obtained properly combining with ABE 98D. The quoted systematic error includes an uncertainty of  $\pm 0.00012$  due to the uncertainty on  $R_C$ .

<sup>2</sup> ACCIARRI 00 obtain this result using a double-tagging technique, with a high  $p_T$  lepton tag and an impact parameter tag in opposite hemispheres.

<sup>3</sup> ABBIENDI 99B tag  $Z \rightarrow b\bar{b}$  decays using leptons and/or separated decay vertices. The  $b$ -tagging efficiency is measured directly from the data using a double-tagging technique.

<sup>4</sup> ABREU 99B obtain this result combining in a multivariate analysis several tagging methods (impact parameter and secondary vertex reconstruction, complemented by event shape variables). For  $R_C$  different from its Standard Model value of 0.172,  $R_b$  varies as  $-0.024 \times (R_C - 0.172)$ .

<sup>5</sup> BARATE 97F combine the lifetime-mass hemisphere tag (BARATE 97E) with event shape information and lepton tag to identify  $Z \rightarrow b\bar{b}$  candidates. They further use  $c$ - and  $u$  $d$ -selection tags to identify the background. For  $R_C$  different from its Standard Model value of 0.172,  $R_b$  varies as  $-0.019 \times (R_C - 0.172)$ .

<sup>6</sup> ABREU 95D perform a maximum likelihood fit to the combined  $p$  and  $p_T$  distributions of single and dilepton samples. The second error includes an uncertainty of  $\pm 0.0023$  due to models and branching ratios.

<sup>7</sup> BUSKULIC 94G perform a simultaneous fit to the  $p$  and  $p_T$  spectra of both single and dilepton events.

<sup>8</sup> JACOBSEN 91 tagged  $b\bar{b}$  events by requiring coincidence of  $\geq 3$  tracks with significant impact parameters using vertex detector. Systematic error includes lifetime and decay uncertainties ( $\pm 0.014$ ).

 $\Gamma(b\bar{b}b\bar{b})/\Gamma(\text{hadrons})$  $\Gamma_{12}/\Gamma_7$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>5.2 ± 1.9 OUR AVERAGE</b>			
3.6 ± 1.7 ± 2.7	<sup>1</sup> ABBIENDI	01G	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
6.0 ± 1.9 ± 1.4	<sup>2</sup> ABREU	99u	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
<sup>1</sup> ABBIENDI 01G use a sample of four-jet events from hadronic Z decays. To enhance the $b\bar{b}b\bar{b}$ signal, at least three of the four jets are required to have a significantly detached secondary vertex.			
<sup>2</sup> ABREU 99u force hadronic Z decays into 3jets to use all the available phase space and require a $b$ tag for every jet. This decay mode includes primary and secondary $4b$ production, e.g. from gluon splitting to $b\bar{b}$ .			

 $\Gamma(g\bar{g}g)/\Gamma(\text{hadrons})$  $\Gamma_{13}/\Gamma_7$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.6 × 10<sup>-2</sup></b>	95	<sup>1</sup> ABREU	96S	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
<sup>1</sup> This branching ratio is slightly dependent on the jet-finder algorithm. The value we quote is obtained using the JADE algorithm, while using the DURHAM algorithm ABREU 96S obtain an upper limit of $1.5 \times 10^{-2}$ .				

 $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$  $\Gamma_{14}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.01 × 10<sup>-5</sup></b>	95	AALTONEN	14E	CDF $E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
< 5.2 × 10 <sup>-5</sup>	95	<sup>1</sup> ACCIARRI	95G	L3 $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 5.5 × 10 <sup>-5</sup>	95	ABREU	94B	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 2.1 × 10 <sup>-4</sup>	95	DECAMP	92	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 1.4 × 10 <sup>-4</sup>	95	AKRAWY	91F	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> This limit is for both decay modes  $Z \rightarrow \pi^0 \gamma/\gamma\gamma$  which are indistinguishable in ACCIARRI 95G.

 $\Gamma(\eta\gamma)/\Gamma_{\text{total}}$  $\Gamma_{15}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< 7.6 × 10 <sup>-5</sup>	95	ACCIARRI	95G	L3 $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 8.0 × 10 <sup>-5</sup>	95	ABREU	94B	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
<b>&lt; 5.1 × 10<sup>-5</sup></b>	95	DECAMP	92	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 2.0 × 10 <sup>-4</sup>	95	AKRAWY	91F	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

 $\Gamma(\omega\gamma)/\Gamma_{\text{total}}$  $\Gamma_{16}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 6.5 × 10<sup>-4</sup></b>	95	ABREU	94B	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

 $\Gamma(\eta'(958)\gamma)/\Gamma_{\text{total}}$  $\Gamma_{17}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 4.2 × 10<sup>-5</sup></b>	95	DECAMP	92	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

 $\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  $\Gamma_{18}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.46 × 10<sup>-5</sup></b>	95	AALTONEN	14E	CDF $E_{\text{cm}}^{\text{pp}} = 1.96$ TeV
< 5.2 × 10 <sup>-5</sup>	95	<sup>1</sup> ACCIARRI	95G	L3 $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 5.5 × 10 <sup>-5</sup>	95	ABREU	94B	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
< 1.4 × 10 <sup>-4</sup>	95	AKRAWY	91F	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> This limit is for both decay modes  $Z \rightarrow \pi^0 \gamma/\gamma\gamma$  which are indistinguishable in ACCIARRI 95G.

# Gauge & Higgs Boson Particle Listings

## Z

$\Gamma(\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.52 \times 10^{-5}$	95	AALTONEN	14E	CDF	$E_{\text{cm}}^{pp} = 1.96 \text{ TeV}$

$\Gamma(\gamma\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.0 \times 10^{-5}$	95	<sup>1</sup> ACCIARRI	95C	L3	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$<1.7 \times 10^{-5}$	95	<sup>1</sup> ABREU	94B	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$<6.6 \times 10^{-5}$	95	AKRAWY	91F	OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> Limit derived in the context of composite Z model.

$\Gamma(\pi^\pm W^\mp)/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$
The value is for the sum of the charge states indicated.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<7 \times 10^{-5}$	95	DECAMP	92	ALEP	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

$\Gamma(\rho^\pm W^\mp)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
The value is for the sum of the charge states indicated.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<8.3 \times 10^{-5}$	95	DECAMP	92	ALEP	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

$\Gamma(J/\psi(1S)X)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.51^{+0.23}_{-0.25}</math> OUR AVERAGE</b>		Error includes scale factor of 1.1.			
$3.21 \pm 0.21 \pm^{0.19}_{0.28}$	553	<sup>1</sup> ACCIARRI	99F	L3	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$3.9 \pm 0.2 \pm 0.3$	511	<sup>2</sup> ALEXANDER	96B	OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$3.73 \pm 0.39 \pm 0.36$	153	<sup>3</sup> ABREU	94P	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

- <sup>1</sup> ACCIARRI 99F combine  $\mu^+\mu^-$  and  $e^+e^- J/\psi(1S)$  decay channels. The branching ratio for prompt  $J/\psi(1S)$  production is measured to be  $(2.1 \pm 0.6 \pm 0.4^{+0.4}_{-0.2}(\text{theor.})) \times 10^{-4}$ .
- <sup>2</sup> ALEXANDER 96B identify  $J/\psi(1S)$  from the decays into lepton pairs.  $(4.8 \pm 2.4)\%$  of this branching ratio is due to prompt  $J/\psi(1S)$  production (ALEXANDER 96N).
- <sup>3</sup> Combining  $\mu^+\mu^-$  and  $e^+e^-$  channels and taking into account the common systematic errors.  $(7.7^{+6.3}_{-5.4})\%$  of this branching ratio is due to prompt  $J/\psi(1S)$  production.

$\Gamma(J/\psi(1S)\gamma)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.6 \times 10^{-6}$	95	<sup>1</sup> AAD	15i	ATLS	$E_{\text{cm}}^{pp} = 8 \text{ TeV}$

<sup>1</sup> AAD 15i use events with the highest  $p_T$  muon in the pair required to have  $p_T > 20 \text{ GeV}$ , the dimuon mass required to be within 0.2 GeV of the  $J/\psi(1S)$  mass and it's transverse momentum required to be  $> 36 \text{ GeV}$ . The photon is also required to have it's  $p_T > 36 \text{ GeV}$ .

$\Gamma(\psi(2S)X)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.60 \pm 0.29</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.			
$1.6 \pm 0.5 \pm 0.3$	39	<sup>1</sup> ACCIARRI	97J	L3	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$1.6 \pm 0.3 \pm 0.2$	46.9	<sup>2</sup> ALEXANDER	96B	OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$1.60 \pm 0.73 \pm 0.33$	5.4	<sup>3</sup> ABREU	94P	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

- <sup>1</sup> ACCIARRI 97J measure this branching ratio via the decay channel  $\psi(2S) \rightarrow \ell^+\ell^- (\ell = \mu, e)$ .
- <sup>2</sup> ALEXANDER 96B measure this branching ratio via the decay channel  $\psi(2S) \rightarrow J/\psi\pi^+\pi^-$ , with  $J/\psi \rightarrow \ell^+\ell^-$ .
- <sup>3</sup> ABREU 94P measure this branching ratio via decay channel  $\psi(2S) \rightarrow J/\psi\pi^+\pi^-$ , with  $J/\psi \rightarrow \mu^+\mu^-$ .

$\Gamma(\chi_{c1}(1P)X)/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.9 \pm 0.7</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.			
$2.7 \pm 0.6 \pm 0.5$	33	<sup>1</sup> ACCIARRI	97J	L3	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$5.0 \pm 2.1 \pm^{1.5}_{0.9}$	6.4	<sup>2</sup> ABREU	94P	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

- <sup>1</sup> ACCIARRI 97J measure this branching ratio via the decay channel  $\chi_{c1} \rightarrow J/\psi + \gamma$ , with  $J/\psi \rightarrow \ell^+\ell^- (\ell = \mu, e)$ . The  $M(\ell^+\ell^-\gamma) - M(\ell^+\ell^-)$  mass difference spectrum is fitted with two gaussian shapes for  $\chi_{c1}$  and  $\chi_{c2}$ .
- <sup>2</sup> This branching ratio is measured via the decay channel  $\chi_{c1} \rightarrow J/\psi + \gamma$ , with  $J/\psi \rightarrow \mu^+\mu^-$ .

$\Gamma(\chi_{c2}(1P)X)/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.2 \times 10^{-3}$	90	<sup>1</sup> ACCIARRI	97J	L3	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> ACCIARRI 97J derive this limit via the decay channel  $\chi_{c2} \rightarrow J/\psi + \gamma$ , with  $J/\psi \rightarrow \ell^+\ell^- (\ell = \mu, e)$ . The  $M(\ell^+\ell^-\gamma) - M(\ell^+\ell^-)$  mass difference spectrum is fitted with two gaussian shapes for  $\chi_{c1}$  and  $\chi_{c2}$ .

$\Gamma(\Upsilon(1S)X + \Upsilon(2S)X + \Upsilon(3S)X)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma = (\Gamma_{29} + \Gamma_{30} + \Gamma_{31})/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.0 \pm 0.4 \pm 0.22</math></b>	6.4	<sup>1</sup> ALEXANDER	96F	OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> ALEXANDER 96F identify the  $\Upsilon$  (which refers to any of the three lowest bound states) through its decay into  $e^+e^-$  and  $\mu^+\mu^-$ . The systematic error includes an uncertainty of  $\pm 0.2$  due to the production mechanism.

$\Gamma(\Upsilon(1S)X)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.4 \times 10^{-6}$	95	<sup>1</sup> AAD	15i	ATLS	$E_{\text{cm}}^{pp} = 8 \text{ TeV}$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- $<4.4 \times 10^{-5}$  95 <sup>2</sup> ACCIARRI 99F L3  $E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> AAD 15i use events with the highest  $p_T$  muon in the pair required to have  $p_T > 20 \text{ GeV}$ , the dimuon mass required to be in the range 8–12 GeV and it's transverse momentum required to be  $> 36 \text{ GeV}$ . The photon is also required to have it's  $p_T > 36 \text{ GeV}$ .

<sup>2</sup> ACCIARRI 99F search for  $\Upsilon(1S)$  through its decay into  $\ell^+\ell^- (\ell = e \text{ or } \mu)$ .

$\Gamma(\Upsilon(2S)X)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 6.5 \times 10^{-6}$	95	<sup>1</sup> AAD	15i	ATLS	$E_{\text{cm}}^{pp} = 8 \text{ TeV}$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- $<13.9 \times 10^{-5}$  95 <sup>2</sup> ACCIARRI 97R L3  $E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> AAD 15i use events with the highest  $p_T$  muon in the pair required to have  $p_T > 20 \text{ GeV}$ , the dimuon mass required to be in the range 8–12 GeV and it's transverse momentum required to be  $> 36 \text{ GeV}$ . The photon is also required to have it's  $p_T > 36 \text{ GeV}$ .

<sup>2</sup> ACCIARRI 97R search for  $\Upsilon(2S)$  through its decay into  $\ell^+\ell^- (\ell = e \text{ or } \mu)$ .

$\Gamma(\Upsilon(3S)X)/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<5.4 \times 10^{-6}$	95	<sup>1</sup> AAD	15i	ATLS	$E_{\text{cm}}^{pp} = 8 \text{ TeV}$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- $<9.4 \times 10^{-5}$  95 <sup>2</sup> ACCIARRI 97R L3  $E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> AAD 15i use events with the highest  $p_T$  muon in the pair required to have  $p_T > 20 \text{ GeV}$ , the dimuon mass required to be in the range 8–12 GeV and it's transverse momentum required to be  $> 36 \text{ GeV}$ . The photon is also required to have it's  $p_T > 36 \text{ GeV}$ .

<sup>2</sup> ACCIARRI 97R search for  $\Upsilon(3S)$  through its decay into  $\ell^+\ell^- (\ell = e \text{ or } \mu)$ .

$\Gamma((D^0/\bar{D}^0)X)/\Gamma(\text{hadrons})$					$\Gamma_{32}/\Gamma_7$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.296 \pm 0.019 \pm 0.021</math></b>	369	<sup>1</sup> ABREU	93i	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> The  $(D^0/\bar{D}^0)$  states in ABREU 93i are detected by the  $K\pi$  decay mode. This is a corrected result (see the erratum of ABREU 93i).

$\Gamma(D^\pm X)/\Gamma(\text{hadrons})$					$\Gamma_{33}/\Gamma_7$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.174 \pm 0.016 \pm 0.018</math></b>	539	<sup>1</sup> ABREU	93i	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> The  $D^\pm$  states in ABREU 93i are detected by the  $K\pi\pi$  decay mode. This is a corrected result (see the erratum of ABREU 93i).

$\Gamma(D^*(2010)^\pm X)/\Gamma(\text{hadrons})$					$\Gamma_{34}/\Gamma_7$
The value is for the sum of the charge states indicated.					
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.163 \pm 0.019</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.			
$0.155 \pm 0.010 \pm 0.013$	358	<sup>1</sup> ABREU	93i	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$
$0.21 \pm 0.04$	362	<sup>2</sup> DECAMP	91J	ALEP	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

- <sup>1</sup>  $D^*(2010)^\pm$  in ABREU 93i are reconstructed from  $D^0\pi^\pm$ , with  $D^0 \rightarrow K^-\pi^+$ . The new CLEO II measurement of  $B(D^{*\pm} \rightarrow D^0\pi^\pm) = (68.1 \pm 1.6)\%$  is used. This is a corrected result (see the erratum of ABREU 93i).
- <sup>2</sup> DECAMP 91J report  $B(D^*(2010)^+ \rightarrow D^0\pi^+) B(D^0 \rightarrow K^-\pi^+) \Gamma(D^*(2010)^\pm X) / \Gamma(\text{hadrons}) = (5.11 \pm 0.34) \times 10^{-3}$ . They obtained the above number assuming  $B(D^0 \rightarrow K^-\pi^+) = (3.62 \pm 0.34 \pm 0.44)\%$  and  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (55 \pm 4)\%$ . We have rescaled their original result of  $0.26 \pm 0.05$  taking into account the new CLEO II branching ratio  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (68.1 \pm 1.6)\%$ .

$\Gamma(D_{s1}(2536)^\pm X)/\Gamma(\text{hadrons})$					$\Gamma_{35}/\Gamma_7$
$D_{s1}(2536)^\pm$ is an expected orbitally-excited state of the $D_s$ meson.					
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.52 \pm 0.09 \pm 0.06</math></b>	92	<sup>1</sup> HEISTER	02B	ALEP	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> HEISTER 02B reconstruct this meson in the decay modes  $D_{s1}(2536)^\pm \rightarrow D^{*\pm}K^0$  and  $D_{s1}(2536)^\pm \rightarrow D^{*0}K^\pm$ . The quoted branching ratio assumes that the decay width of the  $D_{s1}(2536)$  is saturated by the two measured decay modes.

$\Gamma(D_{sJ}(2573)^\pm X)/\Gamma(\text{hadrons})$					$\Gamma_{36}/\Gamma_7$
$D_{sJ}(2573)^\pm$ is an expected orbitally-excited state of the $D_s$ meson.					
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.83 \pm 0.29 \pm^{0.07}_{0.13}</math></b>	64	<sup>1</sup> HEISTER	02B	ALEP	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> HEISTER 02B reconstruct this meson in the decay mode  $D_{s2}^*(2573)^\pm \rightarrow D^0K^\pm$ . The quoted branching ratio assumes that the detected decay mode represents 45% of the full decay width.

$\Gamma(D^{*'}(2629)^\pm X)/\Gamma(\text{hadrons})$					$\Gamma_{37}/\Gamma_7$
$D^{*'}(2629)^\pm$ is a predicted radial excitation of the $D^*(2010)^\pm$ meson.					
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>searched for</b>	<sup>1</sup> ABBIENDI	01N	OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94 \text{ GeV}$	

<sup>1</sup> ABBIENDI 01N searched for the decay mode  $D^{*'}(2629)^\pm \rightarrow D^{*\pm}\pi^+\pi^-$  with  $D^{*+} \rightarrow D^0\pi^+$ , and  $D^0 \rightarrow K^-\pi^+$ . They quote a 95% CL limit for  $Z \rightarrow D^{*'}(2629)^\pm \times B(D^{*'}(2629)^+ \rightarrow D^{*+}\pi^+\pi^-) < 3.1 \times 10^{-3}$ .

$\Gamma(B^*X)/[\Gamma(BX) + \Gamma(B^*X)]$  $\Gamma_{39}/(\Gamma_{38} + \Gamma_{39})$ 

As the experiments assume different values of the  $b$ -baryon contribution, our average should be taken with caution.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.75 ± 0.04 OUR AVERAGE</b>				
0.760 ± 0.036 ± 0.083		<sup>1</sup> ACKERSTAFF 97M	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.771 ± 0.026 ± 0.070		<sup>2</sup> BUSKULIC 96D	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.72 ± 0.03 ± 0.06		<sup>3</sup> ABREU 95R	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.76 ± 0.08 ± 0.06	1378	<sup>4</sup> ACCIARRI 95B	L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ACKERSTAFF 97M use an inclusive  $B$  reconstruction method and assume a  $(13.2 \pm 4.1)\%$   $b$ -baryon contribution. The value refers to a  $b$ -flavored meson mixture of  $B_U$ ,  $B_D$ , and  $B_S$ .

<sup>2</sup> BUSKULIC 96D use an inclusive reconstruction of  $B$  hadrons and assume a  $(12.2 \pm 4.3)\%$   $b$ -baryon contribution. The value refers to a  $b$ -flavored mixture of  $B_U$ ,  $B_D$ , and  $B_S$ .

<sup>3</sup> ABREU 95R use an inclusive  $B$ -reconstruction method and assume a  $(10 \pm 4)\%$   $b$ -baryon contribution. The value refers to a  $b$ -flavored meson mixture of  $B_U$ ,  $B_D$ , and  $B_S$ .

<sup>4</sup> ACCIARRI 95B assume a  $9.4\%$   $b$ -baryon contribution. The value refers to a  $b$ -flavored mixture of  $B_U$ ,  $B_D$ , and  $B_S$ .

 $\Gamma(B^+X)/\Gamma(\text{hadrons})$  $\Gamma_{40}/\Gamma_7$ 

“OUR EVALUATION” is obtained using our current values for  $f(\bar{b} \rightarrow B^+)$  and  $R_b = \Gamma(b\bar{b})/\Gamma(\text{hadrons})$ . We calculate  $\Gamma(B^+X)/\Gamma(\text{hadrons}) = R_b \times f(\bar{b} \rightarrow B^+)$ . The decay fraction  $f(\bar{b} \rightarrow B^+)$  was provided by the Heavy Flavor Averaging Group (HFAG, [http://www.slac.stanford.edu/xorg/hfag/osc/PDG\\_2009/#FRACZ](http://www.slac.stanford.edu/xorg/hfag/osc/PDG_2009/#FRACZ)).

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0869 ± 0.0019 OUR EVALUATION</b>			
<b>0.0887 ± 0.0030</b>	<sup>1</sup> ABDALLAH 03K	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ABDALLAH 03K measure the production fraction of  $B^+$  mesons in hadronic  $Z$  decays  $f(B^+) = (40.99 \pm 0.82 \pm 1.11)\%$ . The value quoted here is obtained multiplying this production fraction by our value of  $R_b = \Gamma(\bar{b}b)/\Gamma(\text{hadrons})$ .

 $\Gamma(B_S^0X)/\Gamma(\text{hadrons})$  $\Gamma_{41}/\Gamma_7$ 

“OUR EVALUATION” is obtained using our current values for  $f(\bar{b} \rightarrow B_S^0)$  and  $R_b = \Gamma(b\bar{b})/\Gamma(\text{hadrons})$ . We calculate  $\Gamma(B_S^0X)/\Gamma(\text{hadrons}) = R_b \times f(\bar{b} \rightarrow B_S^0)$ . The decay fraction  $f(\bar{b} \rightarrow B_S^0)$  was provided by the Heavy Flavor Averaging Group (HFAG, [http://www.slac.stanford.edu/xorg/hfag/osc/PDG\\_2009/#FRACZ](http://www.slac.stanford.edu/xorg/hfag/osc/PDG_2009/#FRACZ)).

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0227 ± 0.0019 OUR EVALUATION</b>			
seen	<sup>1</sup> ABREU 92M	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
seen	<sup>2</sup> ACTON 92N	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
seen	<sup>3</sup> BUSKULIC 92E	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ABREU 92M reported value is  $\Gamma(B_S^0X) \times B(B_S^0 \rightarrow D_S \mu \nu_\mu X) \times B(D_S \rightarrow \phi \pi)/\Gamma(\text{hadrons}) = (18 \pm 8) \times 10^{-5}$ .

<sup>2</sup> ACTON 92N find evidence for  $B_S^0$  production using  $D_S$ - $\ell$  correlations, with  $D_S^+ \rightarrow \phi \pi^+$  and  $K^*(892) K^+$ . Assuming  $R_b$  from the Standard Model and averaging over the  $e$  and  $\mu$  channels, authors measure the product branching fraction to be  $f(\bar{b} \rightarrow B_S^0) \times B(B_S^0 \rightarrow D_S^+ \ell^+ \nu_\ell X) \times B(D_S^- \rightarrow \phi \pi^-) = (3.9 \pm 1.1 \pm 0.8) \times 10^{-4}$ .

<sup>3</sup> BUSKULIC 92E find evidence for  $B_S^0$  production using  $D_S$ - $\ell$  correlations, with  $D_S^+ \rightarrow \phi \pi^+$  and  $K^*(892) K^+$ . Using  $B(D_S^+ \rightarrow \phi \pi^+) = (2.7 \pm 0.7)\%$  and summing up the  $e$  and  $\mu$  channels, the weighted average product branching fraction is measured to be  $B(\bar{b} \rightarrow B_S^0) \times B(B_S^0 \rightarrow D_S^- \ell^+ \nu_\ell X) = 0.040 \pm 0.011 \pm 0.010 - 0.012$ .

 $\Gamma(B_c^+X)/\Gamma(\text{hadrons})$  $\Gamma_{42}/\Gamma_7$ 

VALUE	DOCUMENT ID	TECN	COMMENT
searched for	<sup>1</sup> ACKERSTAFF 98O	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
searched for	<sup>2</sup> ABREU 97E	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
searched for	<sup>3</sup> BARATE 97H	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ACKERSTAFF 98O searched for the decay modes  $B_c \rightarrow J/\psi \pi^+$ ,  $J/\psi a_1^+$ , and  $J/\psi \ell^+ \nu_\ell$ , with  $J/\psi \rightarrow \ell^+ \ell^-$ ,  $\ell = e, \mu$ . The number of candidates (background) for the three decay modes is  $2(0.63 \pm 0.2)$ ,  $0(1.10 \pm 0.22)$ , and  $1(0.82 \pm 0.19)$  respectively. Interpreting the  $2 B_c \rightarrow J/\psi \pi^+$  candidates as signal, they report  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \pi^+)/\Gamma(\text{hadrons}) = (3.8_{-2.4}^{+5.0} \pm 0.5) \times 10^{-5}$ . Interpreted as background, the 90% CL bounds are  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \pi^+)/\Gamma(\text{hadrons}) < 1.06 \times 10^{-4}$ ,  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi a_1^+)/\Gamma(\text{hadrons}) < 5.29 \times 10^{-4}$ ,  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \ell^+ \nu_\ell)/\Gamma(\text{hadrons}) < 6.96 \times 10^{-5}$ .

<sup>2</sup> ABREU 97E searched for the decay modes  $B_c \rightarrow J/\psi \pi^+$ ,  $J/\psi \ell^+ \nu_\ell$ , and  $J/\psi(3\pi)^+$ , with  $J/\psi \rightarrow \ell^+ \ell^-$ ,  $\ell = e, \mu$ . The number of candidates (background) for the three decay modes is  $1(1.7)$ ,  $0(0.3)$ , and  $1(2.3)$  respectively. They report the following 90% CL limits:  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \pi^+)/\Gamma(\text{hadrons}) < (1.05\text{--}0.84) \times 10^{-4}$ ,  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \ell^+ \nu_\ell)/\Gamma(\text{hadrons}) < (5.8\text{--}5.0) \times 10^{-5}$ ,  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi(3\pi)^+)/\Gamma(\text{hadrons}) < 1.75 \times 10^{-4}$ , where the ranges are due to the predicted  $B_c$  lifetime  $(0.4\text{--}1.4)$  ps.

<sup>3</sup> BARATE 97H searched for the decay modes  $B_c \rightarrow J/\psi \pi^+$  and  $J/\psi \ell^+ \nu_\ell$  with  $J/\psi \rightarrow \ell^+ \ell^-$ ,  $\ell = e, \mu$ . The number of candidates (background) for the two decay modes is  $0(0.44)$  and  $2(0.81)$  respectively. They report the following 90% CL limits:  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \pi^+)/\Gamma(\text{hadrons}) < 3.6 \times 10^{-5}$  and  $\Gamma(B_c^+X) \times B(B_c \rightarrow J/\psi \ell^+ \nu_\ell)/\Gamma(\text{hadrons}) < 5.2 \times 10^{-5}$ .

 $\Gamma(\Lambda_c^+X)/\Gamma(\text{hadrons})$  $\Gamma_{43}/\Gamma_7$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.022 ± 0.005 OUR AVERAGE</b>			
0.024 ± 0.005 ± 0.006	<sup>1</sup> ALEXANDER 96R	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
0.021 ± 0.003 ± 0.005	<sup>2</sup> BUSKULIC 96Y	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ALEXANDER 96R measure  $R_b \times f(b \rightarrow \Lambda_c^+ X) \times B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (0.122 \pm 0.023 \pm 0.010)\%$  in hadronic  $Z$  decays; the value quoted here is obtained using our best value  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3)\%$ . The first error is the total experiment's error and the second error is the systematic error due to the branching fraction uncertainty.

<sup>2</sup> BUSKULIC 96Y obtain the production fraction of  $\Lambda_c^+$  baryons in hadronic  $Z$  decays  $f(b \rightarrow \Lambda_c^+ X) = 0.110 \pm 0.014 \pm 0.006$  using  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (4.4 \pm 0.6)\%$ ; we have rescaled using our best value  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3)\%$  obtaining  $f(b \rightarrow \Lambda_c^+ X) = 0.097 \pm 0.013 \pm 0.025$  where the first error is their total experiment's error and the second error is the systematic error due to the branching fraction uncertainty. The value quoted here is obtained multiplying this production fraction by our value of  $R_b = \Gamma(b\bar{b})/\Gamma(\text{hadrons})$ .

 $\Gamma(\Xi_c^0X)/\Gamma(\text{hadrons})$  $\Gamma_{44}/\Gamma_7$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
seen	<sup>1</sup> ABDALLAH 05C	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ABDALLAH 05C searched for the charmed strange baryon  $\Xi_c^0$  in the decay channel  $\Xi_c^0 \rightarrow \Xi^- \pi^+$  ( $\Xi^- \rightarrow \Lambda \pi^-$ ). The production rate is measured to be  $f_{\Xi_c^0} \times B(\Xi_c^0 \rightarrow \Xi^- \pi^+) = (4.7 \pm 1.4 \pm 1.1) \times 10^{-4}$  per hadronic  $Z$  decay.

 $\Gamma(\Xi_b X)/\Gamma(\text{hadrons})$  $\Gamma_{45}/\Gamma_7$ 

Here  $\Xi_b$  is used as a notation for the strange  $b$ -baryon states  $\Xi_b^-$  and  $\Xi_b^0$ .

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
seen	<sup>1</sup> ABDALLAH 05C	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
seen	<sup>2</sup> BUSKULIC 96T	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
seen	<sup>3</sup> ABREU 95V	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> ABDALLAH 05C searched for the beauty strange baryon  $\Xi_b$  in the inclusive semileptonic decay channel  $\Xi_b \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X$ . Evidence for the  $\Xi_b$  production is seen from the observation of  $\Xi^\mp$  production accompanied by a lepton of the same sign. From the excess of “right-sign” pairs  $\Xi^\mp \ell^\mp$  compared to “wrong-sign” pairs  $\Xi^\mp \ell^\pm$  the production rate is measured to be  $B(b \rightarrow \Xi_b) \times B(\Xi_b \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X) = (3.0 \pm 1.0 \pm 0.3) \times 10^{-4}$  per lepton species, averaged over electrons and muons.

<sup>2</sup> BUSKULIC 96T investigate  $\Xi$ -lepton correlations and find a significant excess of “right-sign” pairs  $\Xi^\mp \ell^\mp$  compared to “wrong-sign” pairs  $\Xi^\mp \ell^\pm$ . This excess is interpreted as evidence for  $\Xi_b$  semileptonic decay. The measured product branching ratio is  $B(b \rightarrow \Xi_b) \times B(\Xi_b \rightarrow X_c X \ell^- \bar{\nu}_\ell) \times B(X_c \rightarrow \Xi^- X') = (5.4 \pm 1.1 \pm 0.8) \times 10^{-4}$  per lepton species, averaged over electrons and muons, with  $X_c$  a charmed baryon.

<sup>3</sup> ABREU 95V observe an excess of “right-sign” pairs  $\Xi^\mp \ell^\mp$  compared to “wrong-sign” pairs  $\Xi^\mp \ell^\pm$  in jets; this excess is interpreted as evidence for the beauty strange baryon  $\Xi_b$  production, with  $\Xi_b \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X$ . They find that the probability for this signal to come from non  $b$ -baryon decays is less than  $5 \times 10^{-4}$  and that  $\Lambda_b$  decays can account for less than 10% of these events. The  $\Xi_b$  production rate is then measured to be  $B(b \rightarrow \Xi_b) \times B(\Xi_b \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X) = (5.9 \pm 2.1 \pm 1.0) \times 10^{-4}$  per lepton species, averaged over electrons and muons.

 $\Gamma(b\text{-baryon } X)/\Gamma(\text{hadrons})$  $\Gamma_{46}/\Gamma_7$ 

“OUR EVALUATION” is obtained using our current values for  $f(b \rightarrow b\text{-baryon})$  and  $R_b = \Gamma(b\bar{b})/\Gamma(\text{hadrons})$ . We calculate  $\Gamma(b\text{-baryon } X)/\Gamma(\text{hadrons}) = R_b \times f(b \rightarrow b\text{-baryon})$ . The decay fraction  $f(b \rightarrow b\text{-baryon})$  was provided by the Heavy Flavor Averaging Group (HFAG, [http://www.slac.stanford.edu/xorg/hfag/osc/PDG\\_2009](http://www.slac.stanford.edu/xorg/hfag/osc/PDG_2009)).

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0197 ± 0.0032 OUR EVALUATION</b>			
<b>0.0221 ± 0.0015 ± 0.0058</b>	<sup>1</sup> BARATE 98V	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> BARATE 98V use the overall number of identified protons in  $b$ -hadron decays to measure  $f(b \rightarrow b\text{-baryon}) = 0.102 \pm 0.007 \pm 0.027$ . They assume  $B(b\text{-baryon} \rightarrow pX) = (58 \pm 6)\%$  and  $\text{BR}(B_S^0 \rightarrow pX) = (8.0 \pm 4.0)\%$ . The value quoted here is obtained multiplying this production fraction by our value of  $R_b = \Gamma(b\bar{b})/\Gamma(\text{hadrons})$ .

 $\Gamma(\text{anomalous } \gamma + \text{hadrons})/\Gamma_{\text{total}}$  $\Gamma_{47}/\Gamma$ 

Limits on additional sources of prompt photons beyond expectations for final-state bremsstrahlung.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.2 × 10<sup>-3</sup></b>	95	<sup>1</sup> AKRAWY 90J	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

<sup>1</sup> AKRAWY 90J report  $\Gamma(\gamma X) < 8.2$  MeV at 95%CL. They assume a three-body  $\gamma q \bar{q}$  distribution and use  $E(\gamma) > 10$  GeV.

 $\Gamma(e^+ e^- \gamma)/\Gamma_{\text{total}}$  $\Gamma_{48}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 5.2 × 10<sup>-4</sup></b>	95	<sup>1</sup> ACTON 91B	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV

<sup>1</sup> ACTON 91B looked for isolated photons with  $E > 2\%$  of beam energy ( $> 0.9$  GeV).

 $\Gamma(\mu^+ \mu^- \gamma)/\Gamma_{\text{total}}$  $\Gamma_{49}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 5.6 × 10<sup>-4</sup></b>	95	<sup>1</sup> ACTON 91B	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV

<sup>1</sup> ACTON 91B looked for isolated photons with  $E > 2\%$  of beam energy ( $> 0.9$  GeV).

# Gauge & Higgs Boson Particle Listings

## Z

$\Gamma(\tau^+\tau^-\gamma)/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<7.3 \times 10^{-4}$	95	<sup>1</sup> ACTON 91B	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$	
<sup>1</sup> ACTON 91B looked for isolated photons with $E > 2\%$ of beam energy ( $> 0.9 \text{ GeV}$ ).					

$\Gamma(\ell^+\ell^-\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{51}/\Gamma$
The value is the sum over $\ell = e, \mu, \tau$ .					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6.8 \times 10^{-6}$	95	<sup>1</sup> ACTON 93E	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
<sup>1</sup> For $m_{\gamma\gamma} = 60 \pm 5 \text{ GeV}$ .					

$\Gamma(q\bar{q}\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{52}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<5.5 \times 10^{-6}$	95	<sup>1</sup> ACTON 93E	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
<sup>1</sup> For $m_{\gamma\gamma} = 60 \pm 5 \text{ GeV}$ .					

$\Gamma(\nu\bar{\nu}\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{53}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.1 \times 10^{-6}$	95	<sup>1</sup> ACTON 93E	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
<sup>1</sup> For $m_{\gamma\gamma} = 60 \pm 5 \text{ GeV}$ .					

$\Gamma(e^\pm\mu^\mp)/\Gamma_{\text{total}}$					$\Gamma_{54}/\Gamma$
Test of lepton family number conservation. The value is for the sum of the charge states indicated.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<7.5 \times 10^{-7}$	95	AAD 14AU	ATLS	$E_{\text{cm}}^{p\bar{p}} = 8 \text{ TeV}$	
$<2.5 \times 10^{-6}$	95	ABREU 97c	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<1.7 \times 10^{-6}$	95	AKERS 95W	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<0.6 \times 10^{-5}$	95	ADRIANI 93I	L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<2.6 \times 10^{-5}$	95	DECAMP 92	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	

$\Gamma(e^\pm\mu^\mp)/\Gamma(e^+e^-)$					$\Gamma_{54}/\Gamma_1$
Test of lepton family number conservation. The value is for the sum of the charge states indicated.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.07$	90	ALBAJAR 89	UA1	$E_{\text{cm}}^{p\bar{p}} = 546,630 \text{ GeV}$	

$\Gamma(e^\pm\tau^\mp)/\Gamma_{\text{total}}$					$\Gamma_{55}/\Gamma$
Test of lepton family number conservation. The value is for the sum of the charge states indicated.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.2 \times 10^{-5}$	95	ABREU 97c	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<9.8 \times 10^{-6}$	95	AKERS 95W	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<1.3 \times 10^{-5}$	95	ADRIANI 93I	L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<1.2 \times 10^{-4}$	95	DECAMP 92	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	

$\Gamma(\mu^\pm\tau^\mp)/\Gamma_{\text{total}}$					$\Gamma_{56}/\Gamma$
Test of lepton family number conservation. The value is for the sum of the charge states indicated.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.2 \times 10^{-5}$	95	ABREU 97c	DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<1.7 \times 10^{-5}$	95	AKERS 95W	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<1.9 \times 10^{-5}$	95	ADRIANI 93I	L3	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
$<1.0 \times 10^{-4}$	95	DECAMP 92	ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	

$\Gamma(p\bar{e})/\Gamma_{\text{total}}$					$\Gamma_{57}/\Gamma$
Test of baryon number and lepton number conservations. Charge conjugate states are implied.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.8 \times 10^{-6}$	95	<sup>1</sup> ABBIENDI 99I	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
<sup>1</sup> ABBIENDI 99I give the 95%CL limit on the partial width $\Gamma(Z^0 \rightarrow p\bar{e}) < 4.6 \text{ KeV}$ and we have transformed it into a branching ratio.					

$\Gamma(p\mu)/\Gamma_{\text{total}}$					$\Gamma_{58}/\Gamma$
Test of baryon number and lepton number conservations. Charge conjugate states are implied.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.8 \times 10^{-6}$	95	<sup>1</sup> ABBIENDI 99I	OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$	
<sup>1</sup> ABBIENDI 99I give the 95%CL limit on the partial width $\Gamma(Z^0 \rightarrow p\mu) < 4.4 \text{ KeV}$ and we have transformed it into a branching ratio.					

### AVERAGE PARTICLE MULTIPLICITIES IN HADRONIC Z DECAY

Summed over particle and antiparticle, when appropriate.

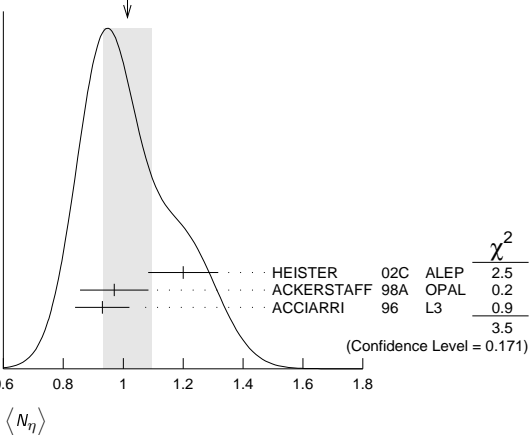
$\langle N_\gamma \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$20.97 \pm 0.02 \pm 1.15$	ACKERSTAFF 98A	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

$\langle N_{\pi^\pm} \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$17.03 \pm 0.16$ OUR AVERAGE			
$17.007 \pm 0.209$	ABE 04C	SLD	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$17.26 \pm 0.10 \pm 0.88$	ABREU 98L	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$17.04 \pm 0.31$	BARATE 98V	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$17.05 \pm 0.43$	AKERS 94P	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

$\langle N_{\pi^0} \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$9.76 \pm 0.26$ OUR AVERAGE			
$9.55 \pm 0.06 \pm 0.75$	ACKERSTAFF 98A	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$9.63 \pm 0.13 \pm 0.63$	BARATE 97J	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$9.90 \pm 0.02 \pm 0.33$	ACCIARRI 96	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$9.2 \pm 0.2 \pm 1.0$	ADAM 96	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

$\langle N_\eta \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$1.01 \pm 0.08$ OUR AVERAGE			Error includes scale factor of 1.3. See the ideogram below.
$1.20 \pm 0.04 \pm 0.11$	HEISTER 02C	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.97 \pm 0.03 \pm 0.11$	ACKERSTAFF 98A	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.93 \pm 0.01 \pm 0.09$	ACCIARRI 96	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

WEIGHTED AVERAGE  
1.01±0.08 (Error scaled by 1.3)



$\langle N_{\rho^\pm} \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$2.57 \pm 0.15$ OUR AVERAGE			
$2.59 \pm 0.03 \pm 0.16$	<sup>1</sup> BEDDALL 09	ALEPH archive,	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$2.40 \pm 0.06 \pm 0.43$	ACKERSTAFF 98A	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

<sup>1</sup> BEDDALL 09 analyse 3.2 million hadronic Z decays as archived by ALEPH collaboration and report a value of  $2.59 \pm 0.03 \pm 0.15 \pm 0.04$ . The first error is statistical, the second systematic, and the third arises from extrapolation to full phase space. We combine the systematic errors in quadrature.

$\langle N_{\rho^0} \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$1.24 \pm 0.10$ OUR AVERAGE			Error includes scale factor of 1.1.
$1.19 \pm 0.10$	ABREU 99J	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$1.45 \pm 0.06 \pm 0.20$	BUSKULIC 96H	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

$\langle N_\eta' \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$1.02 \pm 0.06$ OUR AVERAGE			
$1.00 \pm 0.03 \pm 0.06$	HEISTER 02C	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$1.04 \pm 0.04 \pm 0.14$	ACKERSTAFF 98A	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$1.17 \pm 0.09 \pm 0.15$	ACCIARRI 97D	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

$\langle N_{\eta'} \rangle$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.17 \pm 0.05$ OUR AVERAGE			Error includes scale factor of 2.4.
$0.14 \pm 0.01 \pm 0.02$	ACKERSTAFF 98A	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.25 \pm 0.04$	<sup>1</sup> ACCIARRI 97D	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.068 \pm 0.018 \pm 0.016$	<sup>2</sup> BUSKULIC 92D	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

<sup>1</sup> ACCIARRI 97D obtain this value averaging over the two decay channels  $\eta' \rightarrow \pi^+\pi^-\eta$  and  $\eta' \rightarrow \rho^0\gamma$ .

<sup>2</sup> BUSKULIC 92D obtain this value for  $x > 0.1$ .

See key on page 601

## Gauge &amp; Higgs Boson Particle Listings

Z

 $\langle N_{6(980)} \rangle$ 

VALUE

**0.147 ± 0.011 OUR AVERAGE**

0.164 ± 0.021

0.141 ± 0.007 ± 0.011

DOCUMENT ID TECN COMMENT

ABREU 99J DLPH  $E_{cm}^{ee} = 91.2$  GeVACKERSTAFF 98Q OPAL  $E_{cm}^{ee} = 91.2$  GeV $\langle N_{\pi_0(980)} \rangle$ 

VALUE

**0.27 ± 0.04 ± 0.10**

DOCUMENT ID TECN COMMENT

ACKERSTAFF 98A OPAL  $E_{cm}^{ee} = 91.2$  GeV $\langle N_\phi \rangle$ 

VALUE

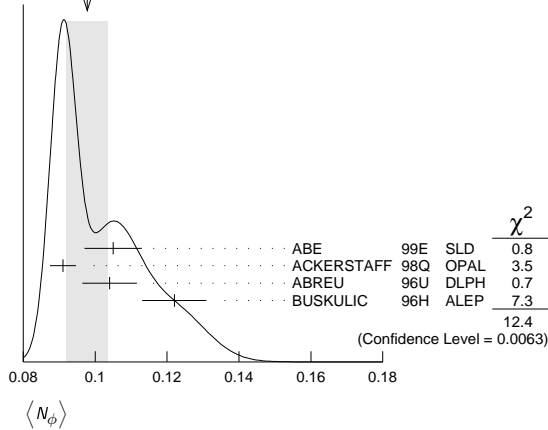
**0.098 ± 0.006 OUR AVERAGE**

0.105 ± 0.008

0.091 ± 0.002 ± 0.003

0.104 ± 0.003 ± 0.007

0.122 ± 0.004 ± 0.008

DOCUMENT ID TECN COMMENT  
Error includes scale factor of 2.0. See the ideogram below.ABE 99E SLD  $E_{cm}^{ee} = 91.2$  GeVACKERSTAFF 98Q OPAL  $E_{cm}^{ee} = 91.2$  GeVABREU 96U DLPH  $E_{cm}^{ee} = 91.2$  GeVBUSKULIC 96H ALEP  $E_{cm}^{ee} = 91.2$  GeVWEIGHTED AVERAGE  
0.098 ± 0.006 (Error scaled by 2.0) $\langle N_{f_2(1270)} \rangle$ 

VALUE

**0.169 ± 0.025 OUR AVERAGE**

0.214 ± 0.038

0.155 ± 0.011 ± 0.018

DOCUMENT ID TECN COMMENT

Error includes scale factor of 1.4.

ABREU 99J DLPH  $E_{cm}^{ee} = 91.2$  GeVACKERSTAFF 98Q OPAL  $E_{cm}^{ee} = 91.2$  GeV $\langle N_{f_1(1285)} \rangle$ 

VALUE

**0.165 ± 0.051**

DOCUMENT ID TECN COMMENT

<sup>1</sup> ABDALLAH 03H DLPH  $E_{cm}^{ee} = 91.2$  GeV<sup>1</sup> ABDALLAH 03H assume a  $K\bar{K}\pi$  branching ratio of  $(9.0 \pm 0.4)\%$ . $\langle N_{f_1(1420)} \rangle$ 

VALUE

**0.056 ± 0.012**

DOCUMENT ID TECN COMMENT

<sup>1</sup> ABDALLAH 03H DLPH  $E_{cm}^{ee} = 91.2$  GeV<sup>1</sup> ABDALLAH 03H assume a  $K\bar{K}\pi$  branching ratio of 100%. $\langle N_{f_2'(1525)} \rangle$ 

VALUE

**0.012 ± 0.006**

DOCUMENT ID TECN COMMENT

ABREU 99J DLPH  $E_{cm}^{ee} = 91.2$  GeV $\langle N_{K^\pm} \rangle$ 

VALUE

**2.24 ± 0.04 OUR AVERAGE**

2.203 ± 0.071

2.21 ± 0.05 ± 0.05

2.26 ± 0.12

2.42 ± 0.13

DOCUMENT ID TECN COMMENT

ABE 04C SLD  $E_{cm}^{ee} = 91.2$  GeVABREU 98L DLPH  $E_{cm}^{ee} = 91.2$  GeVBARATE 98V ALEP  $E_{cm}^{ee} = 91.2$  GeVAKERS 94P OPAL  $E_{cm}^{ee} = 91.2$  GeV $\langle N_{K^0} \rangle$ 

VALUE

**2.039 ± 0.025 OUR AVERAGE**

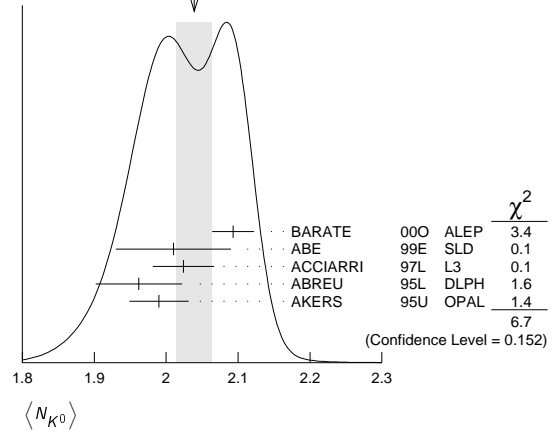
2.093 ± 0.004 ± 0.029

2.01 ± 0.08

2.024 ± 0.006 ± 0.042

1.962 ± 0.022 ± 0.056

1.99 ± 0.01 ± 0.04

DOCUMENT ID TECN COMMENT  
Error includes scale factor of 1.3. See the ideogram below.BARATE 00O ALEP  $E_{cm}^{ee} = 91.2$  GeVABE 99E SLD  $E_{cm}^{ee} = 91.2$  GeVACCIARRI 97L L3  $E_{cm}^{ee} = 91.2$  GeVABREU 95L DLPH  $E_{cm}^{ee} = 91.2$  GeVAKERS 95U OPAL  $E_{cm}^{ee} = 91.2$  GeVWEIGHTED AVERAGE  
2.039 ± 0.025 (Error scaled by 1.3) $\langle N_{K^*(892)} \rangle$ 

VALUE

**0.72 ± 0.05 OUR AVERAGE**

0.712 ± 0.031 ± 0.059

0.72 ± 0.02 ± 0.08

DOCUMENT ID TECN COMMENT

ABREU 95L DLPH  $E_{cm}^{ee} = 91.2$  GeVACTON 93 OPAL  $E_{cm}^{ee} = 91.2$  GeV $\langle N_{K^*(892)^0} \rangle$ 

VALUE

**0.739 ± 0.022 OUR AVERAGE**

0.707 ± 0.041

0.74 ± 0.02 ± 0.02

0.77 ± 0.02 ± 0.07

0.83 ± 0.01 ± 0.09

0.97 ± 0.18 ± 0.31

DOCUMENT ID TECN COMMENT

ABE 99E SLD  $E_{cm}^{ee} = 91.2$  GeVACKERSTAFF 97S OPAL  $E_{cm}^{ee} = 91.2$  GeVABREU 96U DLPH  $E_{cm}^{ee} = 91.2$  GeVBUSKULIC 96H ALEP  $E_{cm}^{ee} = 91.2$  GeVABREU 93 DLPH  $E_{cm}^{ee} = 91.2$  GeV $\langle N_{K_2^*(1430)} \rangle$ 

VALUE

**0.073 ± 0.023**ABREU 99J DLPH  $E_{cm}^{ee} = 91.2$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.19 ± 0.04 ± 0.06

<sup>1</sup> AKERS 95X OPAL  $E_{cm}^{ee} = 91.2$  GeV<sup>1</sup> AKERS 95X obtain this value for  $x < 0.3$ . $\langle N_{D^\pm} \rangle$ 

VALUE

**0.187 ± 0.020 OUR AVERAGE**

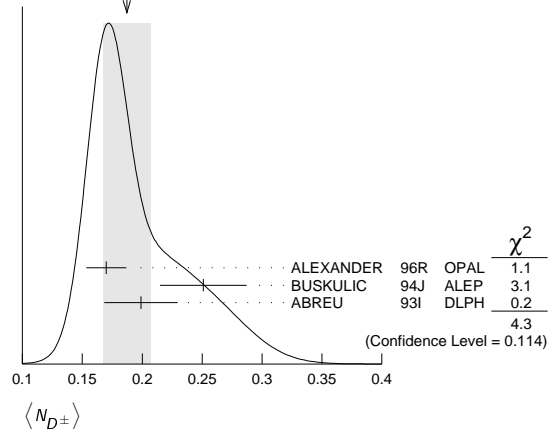
0.170 ± 0.009 ± 0.014

0.251 ± 0.026 ± 0.025

0.199 ± 0.019 ± 0.024

DOCUMENT ID TECN COMMENT

Error includes scale factor of 1.5. See the ideogram below.

ALEXANDER 96R OPAL  $E_{cm}^{ee} = 91.2$  GeVBUSKULIC 94J ALEP  $E_{cm}^{ee} = 91.2$  GeV<sup>1</sup> ABREU 93I DLPH  $E_{cm}^{ee} = 91.2$  GeV<sup>1</sup> See ABREU 95 (erratum).WEIGHTED AVERAGE  
0.187 ± 0.020 (Error scaled by 1.5)

# Gauge & Higgs Boson Particle Listings

## Z

### $\langle N_{D^0} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.462 \pm 0.026</math> OUR AVERAGE</b>			
$0.465 \pm 0.017 \pm 0.027$	ALEXANDER 96R	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.518 \pm 0.052 \pm 0.035$	BUSKULIC 94J	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.403 \pm 0.038 \pm 0.044$	<sup>1</sup> ABREU 93I	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
<sup>1</sup> See ABREU 95 (erratum).			

### $\langle N_{D_s^\pm} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.131 \pm 0.010 \pm 0.018</math></b>	ALEXANDER 96R	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{D^{*}(2010)^\pm} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.183 \pm 0.008</math> OUR AVERAGE</b>			
$0.1854 \pm 0.0041 \pm 0.0091$	<sup>1</sup> ACKERSTAFF 98E	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.187 \pm 0.015 \pm 0.013$	BUSKULIC 94J	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.171 \pm 0.012 \pm 0.016$	<sup>2</sup> ABREU 93I	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
<sup>1</sup> ACKERSTAFF 98E systematic error includes an uncertainty of $\pm 0.0069$ due to the branching ratios $\text{B}(D^{*+} \rightarrow D^0 \pi^+) = 0.683 \pm 0.014$ and $\text{B}(D^0 \rightarrow K^- \pi^+) = 0.0383 \pm 0.0012$ .			
<sup>2</sup> See ABREU 95 (erratum).			

### $\langle N_{D_{s1}(2536)^+} \rangle$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.9^{+0.7}_{-0.6} \pm 0.2$	<sup>1</sup> ACKERSTAFF 97W	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
<sup>1</sup> ACKERSTAFF 97W obtain this value for $x > 0.6$ and with the assumption that its decay width is saturated by the $D^* K$ final states.			

### $\langle N_{B^*} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.28 \pm 0.01 \pm 0.03</math></b>	<sup>1</sup> ABREU 95R	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
<sup>1</sup> ABREU 95R quote this value for a flavor-averaged excited state.			

### $\langle N_{J/\psi(1S)} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0056 \pm 0.0003 \pm 0.0004</math></b>	<sup>1</sup> ALEXANDER 96B	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
<sup>1</sup> ALEXANDER 96B identify $J/\psi(1S)$ from the decays into lepton pairs.			

### $\langle N_{\psi(2S)} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0023 \pm 0.0004 \pm 0.0003</math></b>	ALEXANDER 96B	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_p \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.046 \pm 0.026</math> OUR AVERAGE</b>			
$1.054 \pm 0.035$	ABE 04C	SLD	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$1.08 \pm 0.04 \pm 0.03$	ABREU 98L	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$1.00 \pm 0.07$	BARATE 98V	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.92 \pm 0.11$	AKERS 94P	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

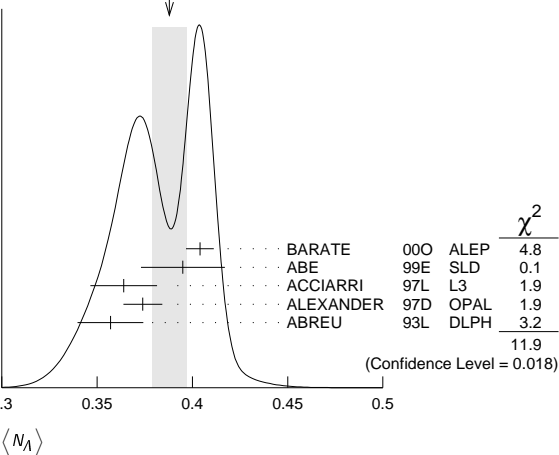
### $\langle N_{\Delta(1232)^{++}} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.087 \pm 0.033</math> OUR AVERAGE</b>	Error includes scale factor of 2.4.		
$0.079 \pm 0.009 \pm 0.011$	ABREU 95W	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.22 \pm 0.04 \pm 0.04$	ALEXANDER 95D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_\Lambda \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.388 \pm 0.009</math> OUR AVERAGE</b>	Error includes scale factor of 1.7. See the ideogram below.		
$0.404 \pm 0.002 \pm 0.007$	BARATE 00O	ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.395 \pm 0.022$	ABE 99E	SLD	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.364 \pm 0.004 \pm 0.017$	ACCIARRI 97L	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.374 \pm 0.002 \pm 0.010$	ALEXANDER 97D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.357 \pm 0.003 \pm 0.017$	ABREU 93L	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

WEIGHTED AVERAGE  
 $0.388 \pm 0.009$  (Error scaled by 1.7)



### $\langle N_{\Lambda(1520)} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0224 \pm 0.0027</math> OUR AVERAGE</b>			
$0.029 \pm 0.005 \pm 0.005$	ABREU 00P	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.0213 \pm 0.0021 \pm 0.0019$	ALEXANDER 97D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Sigma^+} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.107 \pm 0.010</math> OUR AVERAGE</b>			
$0.114 \pm 0.011 \pm 0.009$	ACCIARRI 00J	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.099 \pm 0.008 \pm 0.013$	ALEXANDER 97E	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Sigma^-} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.082 \pm 0.007</math> OUR AVERAGE</b>			
$0.081 \pm 0.002 \pm 0.010$	ABREU 00P	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.083 \pm 0.006 \pm 0.009$	ALEXANDER 97E	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Sigma^+ + \Sigma^-} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.181 \pm 0.018</math> OUR AVERAGE</b>			
$0.182 \pm 0.010 \pm 0.016$	<sup>1</sup> ALEXANDER 97E	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.170 \pm 0.014 \pm 0.061$	ABREU 95O	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

<sup>1</sup> We have combined the values of  $\langle N_{\Sigma^+} \rangle$  and  $\langle N_{\Sigma^-} \rangle$  from ALEXANDER 97E adding the statistical and systematic errors of the two final states separately in quadrature. If isospin symmetry is assumed this value becomes  $0.174 \pm 0.010 \pm 0.015$ .

### $\langle N_{\Sigma^0} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.076 \pm 0.010</math> OUR AVERAGE</b>			
$0.095 \pm 0.015 \pm 0.013$	ACCIARRI 00J	L3	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.071 \pm 0.012 \pm 0.013$	ALEXANDER 97E	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.070 \pm 0.010 \pm 0.010$	ADAM 96B	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{(\Sigma^+ + \Sigma^- + \Sigma^0)/3} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.084 \pm 0.005 \pm 0.008</math></b>	ALEXANDER 97E	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Sigma(1385)^+} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0239 \pm 0.0009 \pm 0.0012</math></b>	ALEXANDER 97D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Sigma(1385)^-} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0240 \pm 0.0010 \pm 0.0014</math></b>	ALEXANDER 97D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Sigma(1385)^+ + \Sigma(1385)^-} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.046 \pm 0.004</math> OUR AVERAGE</b>	Error includes scale factor of 1.6.		
$0.0479 \pm 0.0013 \pm 0.0026$	ALEXANDER 97D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.0382 \pm 0.0028 \pm 0.0045$	ABREU 95O	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

### $\langle N_{\Xi^-} \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0258 \pm 0.0009</math> OUR AVERAGE</b>			
$0.0247 \pm 0.0009 \pm 0.0025$	ABDALLAH 06E	DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$
$0.0259 \pm 0.0004 \pm 0.0009$	ALEXANDER 97D	OPAL	$E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

See key on page 601

## Gauge &amp; Higgs Boson Particle Listings

Z

 $\langle N_{\Xi(1530)^0} \rangle$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0059 ± 0.0011 OUR AVERAGE</b>	Error includes scale factor of 2.3.		
0.0045 ± 0.0005 ± 0.0006	ABDALLAH 05c	DLPH	$E_{cm}^{ee} = 91.2 \text{ GeV}$
0.0068 ± 0.0005 ± 0.0004	ALEXANDER 97D	OPAL	$E_{cm}^{ee} = 91.2 \text{ GeV}$

 $\langle N_{\Omega^-} \rangle$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.00164 ± 0.00028 OUR AVERAGE</b>			
0.0018 ± 0.0003 ± 0.0002	ALEXANDER 97D	OPAL	$E_{cm}^{ee} = 91.2 \text{ GeV}$
0.0014 ± 0.0002 ± 0.0004	ADAM 96B	DLPH	$E_{cm}^{ee} = 91.2 \text{ GeV}$

 $\langle N_{\Lambda_c^+} \rangle$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.078 ± 0.012 ± 0.012</b>	ALEXANDER 96R	OPAL	$E_{cm}^{ee} = 91.2 \text{ GeV}$

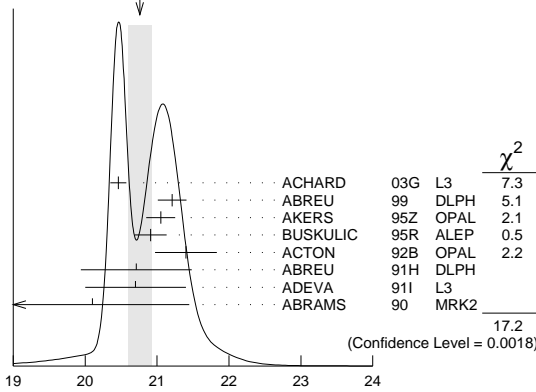
 $\langle N_D^- \rangle$ 

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.078 ± 0.012 ± 0.012</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5.9 ± 1.8 ± 0.5	<sup>1</sup> SCHAEEL 06A	ALEP	$E_{cm}^{ee} = 91.2 \text{ GeV}$

<sup>1</sup>SCHAEEL 06A obtain this anti-deuteron production rate per hadronic Z decay in the anti-deuteron momentum range from 0.62 to 1.03 GeV/c.

 $\langle N_{\text{charged}} \rangle$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>20.76 ± 0.16 OUR AVERAGE</b>	Error includes scale factor of 2.1. See the ideogram below.		
20.46 ± 0.01 ± 0.11	ACHARD 03G	L3	$E_{cm}^{ee} = 91.2 \text{ GeV}$
21.21 ± 0.01 ± 0.20	ABREU 99	DLPH	$E_{cm}^{ee} = 91.2 \text{ GeV}$
21.05 ± 0.20	AKERS 95Z	OPAL	$E_{cm}^{ee} = 91.2 \text{ GeV}$
20.91 ± 0.03 ± 0.22	BUSKULIC 95R	ALEP	$E_{cm}^{ee} = 91.2 \text{ GeV}$
21.40 ± 0.43	ACTON 92B	OPAL	$E_{cm}^{ee} = 91.2 \text{ GeV}$
20.71 ± 0.04 ± 0.77	ABREU 91H	DLPH	$E_{cm}^{ee} = 91.2 \text{ GeV}$
20.7 ± 0.7	ADEVA 91I	L3	$E_{cm}^{ee} = 91.2 \text{ GeV}$
20.1 ± 1.0 ± 0.9	ABRAMS 90	MRK2	$E_{cm}^{ee} = 91.1 \text{ GeV}$

WEIGHTED AVERAGE  
20.76 ± 0.16 (Error scaled by 2.1) $\langle N_{\text{charged}} \rangle$ 

## Z HADRONIC POLE CROSS SECTION

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note "The Z boson" and ref. LEP-SLC 06). This quantity is defined as

$$\sigma_h^0 = \frac{12\pi}{M_Z^2} \frac{\Gamma(e^+e^-)\Gamma(\text{hadrons})}{\Gamma_Z^2}$$

It is one of the parameters used in the Z lineshape fit.

VALUE (nb)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>41.541 ± 0.037 OUR FIT</b>				
41.501 ± 0.055	4.10M	<sup>1</sup> ABBIENDI 01A	OPAL	$E_{cm}^{ee} = 88-94 \text{ GeV}$
41.578 ± 0.069	3.70M	ABREU 00F	DLPH	$E_{cm}^{ee} = 88-94 \text{ GeV}$
41.535 ± 0.055	3.54M	ACCIARRI 00c	L3	$E_{cm}^{ee} = 88-94 \text{ GeV}$
41.559 ± 0.058	4.07M	<sup>2</sup> BARATE 00c	ALEP	$E_{cm}^{ee} = 88-94 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
42 ± 4	45.0	ABRAMS 89B	MRK2	$E_{cm}^{ee} = 89.2-93.0 \text{ GeV}$

<sup>1</sup>ABBIENDI 01A error includes approximately 0.031 due to statistics, 0.033 due to event selection systematics, 0.029 due to uncertainty in luminosity measurement, and 0.011 due to LEP energy uncertainty.

<sup>2</sup>BARATE 00c error includes approximately 0.030 due to statistics, 0.026 due to experimental systematics, and 0.025 due to uncertainty in luminosity measurement.

## Z VECTOR COUPLINGS

These quantities are the effective vector couplings of the Z to charged leptons. Their magnitude is derived from a measurement of the Z lineshape and the forward-backward lepton asymmetries as a function of energy around the Z mass. The relative sign among the vector to axial-vector couplings is obtained from a measurement of the Z asymmetry parameters,  $A_e$ ,  $A_\mu$ , and  $A_\tau$ . By convention the sign of  $g_A^e$  is fixed to be negative (and opposite to that of  $g_V^e$  obtained using  $\nu_e$  scattering measurements). For the light quarks, the sign of the couplings is assigned consistently with this assumption. The fit values quoted below correspond to global nine- or five-parameter fits to lineshape, lepton forward-backward asymmetry, and  $A_e$ ,  $A_\mu$ , and  $A_\tau$  measurements. See the note "The Z boson" and ref. LEP-SLC 06 for details. Where  $p\bar{p}$  and  $e p$  data is quoted, OUR FIT value corresponds to a weighted average of this with the LEP/SLD fit result.

 $g_V^e$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.03817 ± 0.00047 OUR FIT</b>				
-0.058 ± 0.016 ± 0.007	5026	<sup>1</sup> ACOSTA 05M	CDF	$E_{cm}^{p\bar{p}} = 1.96 \text{ TeV}$
-0.0346 ± 0.0023	137.0K	<sup>2</sup> ABBIENDI 01o	OPAL	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0412 ± 0.0027	124.4k	<sup>3</sup> ACCIARRI 00c	L3	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0400 ± 0.0037		BARATE 00c	ALEP	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0414 ± 0.0020		<sup>4</sup> ABE 95J	SLD	$E_{cm}^{ee} = 91.31 \text{ GeV}$

<sup>1</sup>ACOSTA 05M determine the forward-backward asymmetry of  $e^+e^-$  pairs produced via  $q\bar{q} \rightarrow Z/\gamma^* \rightarrow e^+e^-$  in 15  $M(e^+e^-)$  effective mass bins ranging from 40 GeV to 600 GeV. These results are used to obtain the vector and axial-vector couplings of the Z to  $e^+e^-$ , assuming the quark couplings are as predicted by the standard model. Higher order radiative corrections have not been taken into account.

<sup>2</sup>ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>3</sup>ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

<sup>4</sup>ABE 95J obtain this result combining polarized Bhabha results with the  $A_{LR}$  measurement of ABE 94c. The Bhabha results alone give  $-0.0507 \pm 0.0096 \pm 0.0020$ .

 $g_V^u$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.0367 ± 0.0023 OUR FIT</b>				
-0.0388 ± 0.0060 -0.0064	182.8K	<sup>1</sup> ABBIENDI 01o	OPAL	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0386 ± 0.0073	113.4k	<sup>2</sup> ACCIARRI 00c	L3	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0362 ± 0.0061		BARATE 00c	ALEP	$E_{cm}^{ee} = 88-94 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.0413 ± 0.0060	66143	<sup>3</sup> ABBIENDI 01K	OPAL	$E_{cm}^{ee} = 89-93 \text{ GeV}$

<sup>1</sup>ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>2</sup>ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

<sup>3</sup>ABBIENDI 01K obtain this from an angular analysis of the muon pair asymmetry which takes into account effects of initial state radiation on an event by event basis and of initial-final state interference.

 $g_V^d$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.0366 ± 0.0010 OUR FIT</b>				
-0.0365 ± 0.0023	151.5K	<sup>1</sup> ABBIENDI 01o	OPAL	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0384 ± 0.0026	103.0k	<sup>2</sup> ACCIARRI 00c	L3	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0361 ± 0.0068		BARATE 00c	ALEP	$E_{cm}^{ee} = 88-94 \text{ GeV}$

<sup>1</sup>ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>2</sup>ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

 $g_V^s$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.03783 ± 0.00041 OUR FIT</b>				
-0.0358 ± 0.0014	471.3K	<sup>1</sup> ABBIENDI 01o	OPAL	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0397 ± 0.0020	379.4k	<sup>2</sup> ABREU 00F	DLPH	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0397 ± 0.0017	340.8k	<sup>3</sup> ACCIARRI 00c	L3	$E_{cm}^{ee} = 88-94 \text{ GeV}$
-0.0383 ± 0.0018	500k	BARATE 00c	ALEP	$E_{cm}^{ee} = 88-94 \text{ GeV}$

<sup>1</sup>ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>2</sup>Using forward-backward lepton asymmetries.

<sup>3</sup>ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

 $g_V^b$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.25 +0.07 -0.06 OUR AVERAGE</b>				
0.201 ± 0.112	156k	<sup>1</sup> ABAZOV 11D	D0	$E_{cm}^{p\bar{p}} = 1.97 \text{ TeV}$
0.27 ± 0.13	1500	<sup>2</sup> AKTAS 06	H1	$e^\pm p \rightarrow \tau_e(\nu_e) X$ , $\sqrt{s} \approx 300 \text{ GeV}$
0.24 ± 0.28 -0.11		<sup>3</sup> LEP-SLC 06		$E_{cm}^{ee} = 88-94 \text{ GeV}$
0.399 ± 0.152 -0.188	5026	<sup>4</sup> ACOSTA 05M	CDF	$E_{cm}^{p\bar{p}} = 1.96 \text{ TeV}$

# Gauge & Higgs Boson Particle Listings

## Z

<sup>1</sup> ABAZOV 11d study  $p\bar{p} \rightarrow Z/\gamma^* e^+e^-$  events using 5 fb<sup>-1</sup> data at  $\sqrt{s} = 1.96$  TeV. The candidate events are selected by requiring two isolated electromagnetic showers with  $E_T > 25$  GeV, at least one electron in the central region and the di-electron mass in the range 50–1000 GeV. From the forward-backward asymmetry, determined as a function of the di-electron mass, they derive the axial and vector couplings of the  $u$ - and  $d$ -quarks and the value of  $\sin^2\theta_{eff}^e = 0.2309 \pm 0.0008(\text{stat}) \pm 0.0006(\text{syst})$ .

<sup>2</sup> AKTAS 06 fit the neutral current ( $1.5 \leq Q^2 \leq 30,000$  GeV<sup>2</sup>) and charged current ( $1.5 \leq Q^2 \leq 15,000$  GeV<sup>2</sup>) differential cross sections. In the determination of the  $u$ -quark couplings the electron and  $d$ -quark couplings are fixed to their standard model values.

<sup>3</sup> LEP-SLC 06 is a combination of the results from LEP and SLC experiments using light quark tagging.  $s$ - and  $d$ -quark couplings are assumed to be identical.

<sup>4</sup> ACOSTA 05M determine the forward-backward asymmetry of  $e^+e^-$  pairs produced via  $q\bar{q} \rightarrow Z/\gamma^* \rightarrow e^+e^-$  in 15 M( $e^+e^-$ ) effective mass bins ranging from 40 GeV to 600 GeV. These results are used to obtain the vector and axial-vector couplings of the  $Z$  to the light quarks, assuming the electron couplings are as predicted by the Standard Model. Higher order radiative corrections have not been taken into account.

$g_V^d$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.33 \pm_{-0.06}^{+0.05}</math> OUR AVERAGE</b>				
$-0.351 \pm 0.251$	156k	<sup>1</sup> ABAZOV	11D D0	$E_{cm}^{p\bar{p}} = 1.97$ TeV
$-0.33 \pm 0.33$	1500	<sup>2</sup> AKTAS	06 H1	$e^\pm p \rightarrow \mathcal{P}_e(\nu_e)X$ , $\sqrt{s} \approx 300$ GeV
$-0.33 \pm_{-0.07}^{+0.05}$		<sup>3</sup> LEP-SLC	06	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.226 \pm_{-0.290}^{+0.635} \pm 0.090$	5026	<sup>4</sup> ACOSTA	05M CDF	$E_{cm}^{p\bar{p}} = 1.96$ TeV

<sup>1</sup> ABAZOV 11d study  $p\bar{p} \rightarrow Z/\gamma^* e^+e^-$  events using 5 fb<sup>-1</sup> data at  $\sqrt{s} = 1.96$  TeV. The candidate events are selected by requiring two isolated electromagnetic showers with  $E_T > 25$  GeV, at least one electron in the central region and the di-electron mass in the range 50–1000 GeV. From the forward-backward asymmetry, determined as a function of the di-electron mass, they derive the axial and vector couplings of the  $u$ - and  $d$ -quarks and the value of  $\sin^2\theta_{eff}^e = 0.2309 \pm 0.0008(\text{stat}) \pm 0.0006(\text{syst})$ .

<sup>2</sup> AKTAS 06 fit the neutral current ( $1.5 \leq Q^2 \leq 30,000$  GeV<sup>2</sup>) and charged current ( $1.5 \leq Q^2 \leq 15,000$  GeV<sup>2</sup>) differential cross sections. In the determination of the  $d$ -quark couplings the electron and  $u$ -quark couplings are fixed to their standard model values.

<sup>3</sup> LEP-SLC 06 is a combination of the results from LEP and SLC experiments using light quark tagging.  $s$ - and  $d$ -quark couplings are assumed to be identical.

<sup>4</sup> ACOSTA 05M determine the forward-backward asymmetry of  $e^+e^-$  pairs produced via  $q\bar{q} \rightarrow Z/\gamma^* \rightarrow e^+e^-$  in 15 M( $e^+e^-$ ) effective mass bins ranging from 40 GeV to 600 GeV. These results are used to obtain the vector and axial-vector couplings of the  $Z$  to the light quarks, assuming the electron couplings are as predicted by the Standard Model. Higher order radiative corrections have not been taken into account.

## Z AXIAL-VECTOR COUPLINGS

These quantities are the effective axial-vector couplings of the  $Z$  to charged leptons. Their magnitude is derived from a measurement of the  $Z$  lineshape and the forward-backward lepton asymmetries as a function of energy around the  $Z$  mass. The relative sign among the vector to axial-vector couplings is obtained from a measurement of the  $Z$  asymmetry parameters,  $A_e$ ,  $A_\mu$ , and  $A_\tau$ . By convention the sign of  $g_A^e$  is fixed to be negative (and opposite to that of  $g^{f^e}$  obtained using  $\nu_e$  scattering measurements). For the light quarks, the sign of the couplings is assigned consistently with this assumption. The fit values quoted below correspond to global nine- or five-parameter fits to lineshape, lepton forward-backward asymmetry, and  $A_e$ ,  $A_\mu$ , and  $A_\tau$  measurements. See the note “The  $Z$  boson” and ref. LEP-SLC 06 for details. Where  $p\bar{p}$  and  $e p$  data is quoted, OUR FIT value corresponds to a weighted average of this with the LEP/SLD fit result.

$g_A^e$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.50111 \pm 0.00035</math> OUR FIT</b>				
$-0.528 \pm 0.123 \pm 0.059$	5026	<sup>1</sup> ACOSTA	05M CDF	$E_{cm}^{p\bar{p}} = 1.96$ TeV
$-0.50062 \pm 0.00062$	137.0K	<sup>2</sup> ABBIENDI	01o OPAL	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.5015 \pm 0.0007$	124.4k	<sup>3</sup> ACCIARRI	00c L3	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.50166 \pm 0.00057$		BARATE	00c ALEP	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.4977 \pm 0.0045$		<sup>4</sup> ABE	95J SLD	$E_{cm}^{ee} = 91.31$ GeV

<sup>1</sup> ACOSTA 05M determine the forward-backward asymmetry of  $e^+e^-$  pairs produced via  $q\bar{q} \rightarrow Z/\gamma^* \rightarrow e^+e^-$  in 15 M( $e^+e^-$ ) effective mass bins ranging from 40 GeV to 600 GeV. These results are used to obtain the vector and axial-vector couplings of the  $Z$  to  $e^+e^-$ , assuming the quark couplings are as predicted by the standard model. Higher order radiative corrections have not been taken into account.

<sup>2</sup> ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>3</sup> ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

<sup>4</sup> ABE 95J obtain this result combining polarized Bhabha results with the  $A_{LP}$  measurement of ABE 94c. The Bhabha results alone give  $-0.4968 \pm 0.0039 \pm 0.0027$ .

$g_A^u$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.50120 \pm 0.00054</math> OUR FIT</b>				
$-0.50117 \pm 0.00099$	182.8K	<sup>1</sup> ABBIENDI	01o OPAL	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.5009 \pm 0.0014$	113.4k	<sup>2</sup> ACCIARRI	00c L3	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.50046 \pm 0.00093$		BARATE	00c ALEP	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.520 \pm 0.015$	66143	<sup>3</sup> ABBIENDI	01k OPAL	$E_{cm}^{ee} = 89\text{--}93$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>2</sup> ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

<sup>3</sup> ABBIENDI 01k obtain this from an angular analysis of the muon pair asymmetry which takes into account effects of initial state radiation on an event by event basis and of initial-final state interference.

$g_A^T$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.50204 \pm 0.00064</math> OUR FIT</b>				
$-0.50165 \pm 0.00124$	151.5K	<sup>1</sup> ABBIENDI	01o OPAL	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.5023 \pm 0.0017$	103.0k	<sup>2</sup> ACCIARRI	00c L3	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.50216 \pm 0.00100$		BARATE	00c ALEP	$E_{cm}^{ee} = 88\text{--}94$ GeV

<sup>1</sup> ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>2</sup> ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

$g_A^L$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.50123 \pm 0.00026</math> OUR FIT</b>				
$-0.50089 \pm 0.00045$	471.3K	<sup>1</sup> ABBIENDI	01o OPAL	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.5007 \pm 0.0005$	379.4k	ABREU	00f DLPH	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.50153 \pm 0.00053$	340.8k	<sup>2</sup> ACCIARRI	00c L3	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.50150 \pm 0.00046$	500k	BARATE	00c ALEP	$E_{cm}^{ee} = 88\text{--}94$ GeV

<sup>1</sup> ABBIENDI 01o use their measurement of the  $\tau$  polarization in addition to the lineshape and forward-backward lepton asymmetries.

<sup>2</sup> ACCIARRI 00c use their measurement of the  $\tau$  polarization in addition to forward-backward lepton asymmetries.

$g_A^u$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.50 \pm_{-0.06}^{+0.04}</math> OUR AVERAGE</b>				
$0.501 \pm 0.110$	156k	<sup>1</sup> ABAZOV	11D D0	$E_{cm}^{p\bar{p}} = 1.97$ TeV
$0.57 \pm 0.08$	1500	<sup>2</sup> AKTAS	06 H1	$e^\pm p \rightarrow \mathcal{P}_e(\nu_e)X$ , $\sqrt{s} \approx 300$ GeV
$0.47 \pm_{-0.33}^{+0.05}$		<sup>3</sup> LEP-SLC	06	$E_{cm}^{ee} = 88\text{--}94$ GeV
$0.441 \pm_{-0.173}^{+0.207} \pm 0.067$	5026	<sup>4</sup> ACOSTA	05M CDF	$E_{cm}^{p\bar{p}} = 1.96$ TeV

<sup>1</sup> ABAZOV 11d study  $p\bar{p} \rightarrow Z/\gamma^* e^+e^-$  events using 5 fb<sup>-1</sup> data at  $\sqrt{s} = 1.96$  TeV. The candidate events are selected by requiring two isolated electromagnetic showers with  $E_T > 25$  GeV, at least one electron in the central region and the di-electron mass in the range 50–1000 GeV. From the forward-backward asymmetry, determined as a function of the di-electron mass, they derive the axial and vector couplings of the  $u$ - and  $d$ -quarks and the value of  $\sin^2\theta_{eff}^e = 0.2309 \pm 0.0008(\text{stat}) \pm 0.0006(\text{syst})$ .

<sup>2</sup> AKTAS 06 fit the neutral current ( $1.5 \leq Q^2 \leq 30,000$  GeV<sup>2</sup>) and charged current ( $1.5 \leq Q^2 \leq 15,000$  GeV<sup>2</sup>) differential cross sections. In the determination of the  $u$ -quark couplings the electron and  $d$ -quark couplings are fixed to their standard model values.

<sup>3</sup> LEP-SLC 06 is a combination of the results from LEP and SLC experiments using light quark tagging.  $s$ - and  $d$ -quark couplings are assumed to be identical.

<sup>4</sup> ACOSTA 05M determine the forward-backward asymmetry of  $e^+e^-$  pairs produced via  $q\bar{q} \rightarrow Z/\gamma^* \rightarrow e^+e^-$  in 15 M( $e^+e^-$ ) effective mass bins ranging from 40 GeV to 600 GeV. These results are used to obtain the vector and axial-vector couplings of the  $Z$  to the light quarks, assuming the electron couplings are as predicted by the Standard Model. Higher order radiative corrections have not been taken into account.

$g_A^d$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.523 \pm_{-0.029}^{+0.050}</math> OUR AVERAGE</b>				
$-0.497 \pm 0.165$	156k	<sup>1</sup> ABAZOV	11D D0	$E_{cm}^{p\bar{p}} = 1.97$ TeV
$-0.80 \pm 0.24$	1500	<sup>2</sup> AKTAS	06 H1	$e^\pm p \rightarrow \mathcal{P}_e(\nu_e)X$ , $\sqrt{s} \approx 300$ GeV
$-0.52 \pm_{-0.03}^{+0.05}$		<sup>3</sup> LEP-SLC	06	$E_{cm}^{ee} = 88\text{--}94$ GeV
$-0.016 \pm_{-0.536}^{+0.346} \pm 0.091$	5026	<sup>4</sup> ACOSTA	05M CDF	$E_{cm}^{p\bar{p}} = 1.96$ TeV

<sup>1</sup> ABAZOV 11d study  $p\bar{p} \rightarrow Z/\gamma^* e^+e^-$  events using 5 fb<sup>-1</sup> data at  $\sqrt{s} = 1.96$  TeV. The candidate events are selected by requiring two isolated electromagnetic showers with  $E_T > 25$  GeV, at least one electron in the central region and the di-electron mass in the range 50–1000 GeV. From the forward-backward asymmetry, determined as a function of the di-electron mass, they derive the axial and vector couplings of the  $u$ - and  $d$ -quarks and the value of  $\sin^2\theta_{eff}^e = 0.2309 \pm 0.0008(\text{stat}) \pm 0.0006(\text{syst})$ .



<sup>2</sup> AKTAS 06 fit the neutral current ( $1.5 \leq Q^2 \leq 30,000 \text{ GeV}^2$ ) and charged current ( $1.5 \leq Q^2 \leq 15,000 \text{ GeV}^2$ ) differential cross sections. In the determination of the  $d$ -quark couplings the electron and  $u$ -quark couplings are fixed to their standard model values.

<sup>3</sup> LEP-SLC 06 is a combination of the results from LEP and SLC experiments using light quark tagging.  $s$ - and  $d$ -quark couplings are assumed to be identical.

<sup>4</sup> ACOSTA 05M determine the forward-backward asymmetry of  $e^+e^-$  pairs produced via  $q\bar{q} \rightarrow Z/\gamma^* \rightarrow e^+e^-$  in 15  $M(e^+e^-)$  effective mass bins ranging from 40 GeV to 600 GeV. These results are used to obtain the vector and axial-vector couplings of the Z to the light quarks, assuming the electron couplings are as predicted by the Standard Model. Higher order radiative corrections have not been taken into account.

### Z COUPLINGS TO NEUTRAL LEPTONS

Averaging over neutrino species, the invisible Z decay width determines the effective neutrino coupling  $g^{\nu e}$ . For  $g^{\nu e}$  and  $g^{\nu \mu}$ ,  $\nu_e e$  and  $\nu_\mu e$  scattering results are combined with  $g_A^e$  and  $g_V^e$  measurements at the Z mass to obtain  $g^{\nu e}$  and  $g^{\nu \mu}$  following NOVIKOV 93c.

#### $g^{\nu e}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.50076 ± 0.00076</b>	<sup>1</sup> LEP-SLC	06	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> From invisible Z-decay width.

#### $g^{\nu \mu}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.528 ± 0.085</b>	<sup>1</sup> VILAIN	94	CHM2 From $\nu_\mu e$ and $\nu_e e$ scattering

<sup>1</sup> VILAIN 94 derive this value from their value of  $g^{\nu \mu}$  and their ratio  $g^{\nu e}/g^{\nu \mu} = 1.05^{+0.15}_{-0.18}$ .

#### $g^{\nu \mu}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.502 ± 0.017</b>	<sup>1</sup> VILAIN	94	CHM2 From $\nu_\mu e$ scattering

<sup>1</sup> VILAIN 94 derive this value from their measurement of the couplings  $g_A^{\nu \mu} = -0.503 \pm 0.017$  and  $g_V^{\nu \mu} = -0.035 \pm 0.017$  obtained from  $\nu_\mu e$  scattering. We have re-evaluated this value using the current PDG values for  $g_A^e$  and  $g_V^e$ .

### Z ASYMMETRY PARAMETERS

For each fermion-antifermion pair coupling to the Z these quantities are defined as

$$A_f = \frac{2g_V^f g_A^f}{(g_V^f)^2 + (g_A^f)^2}$$

where  $g_V^f$  and  $g_A^f$  are the effective vector and axial-vector couplings. For their relation to the various lepton asymmetries see the note "The Z boson" and ref. LEP-SLC 06.

#### $A_e$

Using polarized beams, this quantity can also be measured as  $(\sigma_L - \sigma_R)/(\sigma_L + \sigma_R)$ , where  $\sigma_L$  and  $\sigma_R$  are the  $e^+e^-$  production cross sections for Z bosons produced with left-handed and right-handed electrons respectively.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1515 ± 0.0019 OUR AVERAGE</b>				
0.1454 ± 0.0108 ± 0.0036	144810	<sup>1</sup> ABBIENDI	01o	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.1516 ± 0.0021	559000	<sup>2</sup> ABE	01B	SLD $E_{\text{cm}}^{\text{ee}} = 91.24 \text{ GeV}$
0.1504 ± 0.0068 ± 0.0008		<sup>3</sup> HEISTER	01	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.1382 ± 0.0116 ± 0.0005	105000	<sup>4</sup> ABREU	00E	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.1678 ± 0.0127 ± 0.0030	137092	<sup>5</sup> ACCIARRI	98H	L3 $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.162 ± 0.041 ± 0.014	89838	<sup>6</sup> ABE	97	SLD $E_{\text{cm}}^{\text{ee}} = 91.27 \text{ GeV}$
0.202 ± 0.038 ± 0.008		<sup>7</sup> ABE	95J	SLD $E_{\text{cm}}^{\text{ee}} = 91.31 \text{ GeV}$

<sup>1</sup> ABBIENDI 01o fit for  $A_e$  and  $A_\tau$  from measurements of the  $\tau$  polarization at varying  $\tau$  production angles. The correlation between  $A_e$  and  $A_\tau$  is less than 0.03.

<sup>2</sup> ABE 01B use the left-right production and left-right forward-backward decay asymmetries in leptonic Z decays to obtain a value of 0.1544 ± 0.0060. This is combined with left-right production asymmetry measurement using hadronic Z decays (ABE 00B) to obtain the quoted value.

<sup>3</sup> HEISTER 01 obtain this result fitting the  $\tau$  polarization as a function of the polar production angle of the  $\tau$ .

<sup>4</sup> ABREU 00E obtain this result fitting the  $\tau$  polarization as a function of the polar  $\tau$  production angle. This measurement is a combination of different analyses (exclusive  $\tau$  decay modes, inclusive hadronic 1-prong reconstruction, and a neural network analysis).

<sup>5</sup> Derived from the measurement of forward-backward  $\tau$  polarization asymmetry.

<sup>6</sup> ABE 97 obtain this result from a measurement of the observed left-right charge asymmetry,  $A_{\text{LR}}^{\text{Obs}} = 0.225 \pm 0.056 \pm 0.019$ , in hadronic Z decays. If they combine this value of  $A_{\text{LR}}^{\text{Obs}}$  with their earlier measurement of  $A_{\text{LR}}^{\text{Obs}}$  they determine  $A_e$  to be  $0.1574 \pm 0.0197 \pm 0.0067$  independent of the beam polarization.

<sup>7</sup> ABE 95J obtain this result from polarized Bhabha scattering.

#### $A_\mu$

This quantity is directly extracted from a measurement of the left-right forward-backward asymmetry in  $\mu^+\mu^-$  production at SLC using a polarized electron beam. This double asymmetry eliminates the dependence on the Z-e-e coupling parameter  $A_e$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.142 ± 0.015</b>	16844	<sup>1</sup> ABE	01B	SLD $E_{\text{cm}}^{\text{ee}} = 91.24 \text{ GeV}$
0.153 ± 0.012	1.7M	<sup>2</sup> AAD	15Bt	ATLS $E_{\text{cm}}^{\text{pp}} = 7 \text{ TeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> ABE 01B obtain this direct measurement using the left-right production and left-right forward-backward polar angle asymmetries in  $\mu^+\mu^-$  decays of the Z boson obtained with a polarized electron beam.

<sup>2</sup> AAD 15Bt study  $pp \rightarrow Z \rightarrow \ell^+\ell^-$  events where  $\ell$  is an electron or a muon in the dilepton mass region 70–1000 GeV. The background in the Z peak region is estimated to be < 1% for the muon channel. The muon asymmetry parameter is derived from the measured forward-backward asymmetry assuming the value of the quark asymmetry parameter from the SM. For this reason it is not used in the average.

#### $A_\tau$

The LEP Collaborations derive this quantity from the measurement of the  $\tau$  polarization in  $Z \rightarrow \tau^+\tau^-$ . The SLD Collaboration directly extracts this quantity from its measured left-right forward-backward asymmetry in  $Z \rightarrow \tau^+\tau^-$  produced using a polarized  $e^-$  beam. This double asymmetry eliminates the dependence on the Z-e-e coupling parameter  $A_e$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.143 ± 0.004 OUR AVERAGE</b>				
0.1456 ± 0.0076 ± 0.0057	144810	<sup>1</sup> ABBIENDI	01o	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.136 ± 0.015	16083	<sup>2</sup> ABE	01B	SLD $E_{\text{cm}}^{\text{ee}} = 91.24 \text{ GeV}$
0.1451 ± 0.0052 ± 0.0029		<sup>3</sup> HEISTER	01	ALEP $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.1359 ± 0.0079 ± 0.0055	105000	<sup>4</sup> ABREU	00E	DLPH $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
0.1476 ± 0.0088 ± 0.0062	137092	ACCIARRI	98H	L3 $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> ABBIENDI 01o fit for  $A_e$  and  $A_\tau$  from measurements of the  $\tau$  polarization at varying  $\tau$  production angles. The correlation between  $A_e$  and  $A_\tau$  is less than 0.03.

<sup>2</sup> ABE 01B obtain this direct measurement using the left-right production and left-right forward-backward polar angle asymmetries in  $\tau^+\tau^-$  decays of the Z boson obtained with a polarized electron beam.

<sup>3</sup> HEISTER 01 obtain this result fitting the  $\tau$  polarization as a function of the polar production angle of the  $\tau$ .

<sup>4</sup> ABREU 00E obtain this result fitting the  $\tau$  polarization as a function of the polar  $\tau$  production angle. This measurement is a combination of different analyses (exclusive  $\tau$  decay modes, inclusive hadronic 1-prong reconstruction, and a neural network analysis).

#### $A_s$

The SLD Collaboration directly extracts this quantity by a simultaneous fit to four measured s-quark polar angle distributions corresponding to two states of  $e^-$  polarization (positive and negative) and to the  $K^+K^-$  and  $K^\pm K_S^0$  strange particle tagging modes in the hadronic final states.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.895 ± 0.066 ± 0.062</b>	2870	<sup>1</sup> ABE	00D	SLD $E_{\text{cm}}^{\text{ee}} = 91.2 \text{ GeV}$

<sup>1</sup> ABE 00D tag  $Z \rightarrow s\bar{s}$  events by an absence of B or D hadrons and the presence in each hemisphere of a high momentum  $K^\pm$  or  $K_S^0$ .

#### $A_c$

This quantity is directly extracted from a measurement of the left-right forward-backward asymmetry in  $c\bar{c}$  production at SLC using polarized electron beam. This double asymmetry eliminates the dependence on the Z-e-e coupling parameter  $A_e$ . OUR FIT is obtained by a simultaneous fit to several c- and b-quark measurements as explained in the note "The Z boson" and ref. LEP-SLC 06.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.670 ± 0.027 OUR FIT</b>			
0.6712 ± 0.0224 ± 0.0157	<sup>1</sup> ABE	05	SLD $E_{\text{cm}}^{\text{ee}} = 91.24 \text{ GeV}$
0.583 ± 0.055 ± 0.055	<sup>2</sup> ABE	02G	SLD $E_{\text{cm}}^{\text{ee}} = 91.24 \text{ GeV}$
0.688 ± 0.041	<sup>3</sup> ABE	01c	SLD $E_{\text{cm}}^{\text{ee}} = 91.25 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> ABE 05 use hadronic Z decays collected during 1996–98 to obtain an enriched sample of  $c\bar{c}$  events tagging on the invariant mass of reconstructed secondary decay vertices. The charge of the underlying c-quark is obtained with an algorithm that takes into account the net charge of the vertex as well as the charge of tracks emanating from the vertex and identified as kaons. This yields (9970 events)  $A_c = 0.6747 \pm 0.0290 \pm 0.0233$ . Taking into account all correlations with earlier results reported in ABE 02G and ABE 01c, they obtain the quoted overall SLD result.

<sup>2</sup> ABE 02G tag b and c quarks through their semileptonic decays into electrons and muons. A maximum likelihood fit is performed to extract simultaneously  $A_b$  and  $A_c$ .

<sup>3</sup> ABE 01c tag  $Z \rightarrow c\bar{c}$  events using two techniques: exclusive reconstruction of  $D^{*+}$ ,  $D^+$  and  $D^0$  mesons and the soft pion tag for  $D^{*+} \rightarrow D^0\pi^+$ . The large background from D mesons produced in  $b\bar{b}$  events is separated efficiently from the signal using precision vertex information. When combining the  $A_c$  values from these two samples, care is taken to avoid double counting of events common to the two samples, and common systematic errors are properly taken into account.

# Gauge & Higgs Boson Particle Listings

## Z

### $A_b$

This quantity is directly extracted from a measurement of the left-right forward-backward asymmetry in  $b\bar{b}$  production at SLC using polarized electron beam. This double asymmetry eliminates the dependence on the  $Z$ - $e$ - $e$  coupling parameter  $A_e$ . OUR FIT is obtained by a simultaneous fit to several  $c$ - and  $b$ -quark measurements as explained in the note "The Z boson" and ref. LEP-SLC 06.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.923 ± 0.020 OUR FIT</b>				
0.9170 ± 0.0147 ± 0.0145		<sup>1</sup> ABE 05	SLD	$E_{\text{cm}}^{\text{ee}} = 91.24$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.907 ± 0.020 ± 0.024	48028	<sup>2</sup> ABE 03f	SLD	$E_{\text{cm}}^{\text{ee}} = 91.24$ GeV
0.919 ± 0.030 ± 0.024		<sup>3</sup> ABE 02g	SLD	$E_{\text{cm}}^{\text{ee}} = 91.24$ GeV
0.855 ± 0.088 ± 0.102	7473	<sup>4</sup> ABE 99L	SLD	$E_{\text{cm}}^{\text{ee}} = 91.27$ GeV

<sup>1</sup> ABE 05 use hadronic Z decays collected during 1996–98 to obtain an enriched sample of  $b\bar{b}$  events tagging on the invariant mass of reconstructed secondary decay vertices. The charge of the underlying  $b$ -quark is obtained with an algorithm that takes into account the net charge of the vertex as well as the charge of tracks emanating from the vertex and identified as kaons. This yields (25917 events)  $A_b = 0.9173 \pm 0.0184 \pm 0.0173$ . Taking into account all correlations with earlier results reported in ABE 03f, ABE 02g and ABE 99L, they obtain the quoted overall SLD result.

<sup>2</sup> ABE 03f obtain an enriched sample of  $b\bar{b}$  events tagging on the invariant mass of a 3-dimensional topologically reconstructed secondary decay. The charge of the underlying  $b$  quark is obtained using a self-calibrating track-charge method. For the 1996–1998 data sample they measure  $A_b = 0.906 \pm 0.022 \pm 0.023$ . The value quoted here is obtained combining the above with the result of ABE 98i (1993–1995 data sample).

<sup>3</sup> ABE 02g tag  $b$  and  $c$  quarks through their semileptonic decays into electrons and muons. A maximum likelihood fit is performed to extract simultaneously  $A_b$  and  $A_c$ .

<sup>4</sup> ABE 99L obtain an enriched sample of  $b\bar{b}$  events tagging with an inclusive vertex mass cut. For distinguishing  $b$  and  $\bar{b}$  quarks they use the charge of identified  $K^\pm$ .

### TRANSVERSE SPIN CORRELATIONS IN $Z \rightarrow \tau^+ \tau^-$

The correlations between the transverse spin components of  $\tau^+ \tau^-$  produced in Z decays may be expressed in terms of the vector and axial-vector couplings:

$$C_{TT} = \frac{|g_A^\tau|^2 - |g_V^\tau|^2}{|g_A^\tau|^2 + |g_V^\tau|^2}$$

$$C_{TN} = -2 \frac{|g_A^\tau| |g_V^\tau|}{|g_A^\tau|^2 + |g_V^\tau|^2} \sin(\Phi_{g_V^\tau} - \Phi_{g_A^\tau})$$

$C_{TT}$  refers to the transverse-transverse (within the collision plane) spin correlation and  $C_{TN}$  refers to the transverse-normal (to the collision plane) spin correlation.

The longitudinal  $\tau$  polarization  $P_\tau$  ( $= -A_\tau$ ) is given by:

$$P_\tau = -2 \frac{|g_A^\tau| |g_V^\tau|}{|g_A^\tau|^2 + |g_V^\tau|^2} \cos(\Phi_{g_V^\tau} - \Phi_{g_A^\tau})$$

Here  $\Phi$  is the phase and the phase difference  $\Phi_{g_V^\tau} - \Phi_{g_A^\tau}$  can be obtained using both the measurements of  $C_{TN}$  and  $P_\tau$ .

### $C_{TT}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.01 ± 0.12 OUR AVERAGE</b>				
0.87 ± 0.20 ± 0.10 – 0.12	9.1k	ABREU	97G DLPH	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV
1.06 ± 0.13 ± 0.05	120k	BARATE	97D ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV

### $C_{TN}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.08 ± 0.13 ± 0.04</b>	120k	<sup>1</sup> BARATE	97D ALEP	$E_{\text{cm}}^{\text{ee}} = 91.2$ GeV

<sup>1</sup> BARATE 97D combine their value of  $C_{TN}$  with the world average  $P_\tau = -0.140 \pm 0.007$  to obtain  $\tan(\Phi_{g_V^\tau} - \Phi_{g_A^\tau}) = -0.57 \pm 0.97$ .

### FORWARD-BACKWARD $e^+ e^- \rightarrow f \bar{f}$ CHARGE ASYMMETRIES

These asymmetries are experimentally determined by tagging the respective lepton or quark flavor in  $e^+ e^-$  interactions. Details of heavy flavor ( $c$ - or  $b$ -quark) tagging at LEP are described in the note on "The Z boson" and ref. LEP-SLC 06. The Standard Model predictions for LEP data have been (re)computed using the ZFITTER package (version 6.36) with input parameters  $M_Z = 91.187$  GeV,  $M_{\text{top}} = 174.3$  GeV,  $M_{\text{Higgs}} = 150$  GeV,  $\alpha_s = 0.119$ ,  $\alpha^{(5)}$  ( $M_Z$ ) =  $1/128.877$  and the Fermi constant  $G_F = 1.16637 \times 10^{-5} \text{ GeV}^{-2}$  (see the note on "The Z boson" for references). For non-LEP data the Standard Model predictions are as given by the authors of the respective publications.

### $A_{\text{FB}}^{(0,e)}$ CHARGE ASYMMETRY IN $e^+ e^- \rightarrow e^+ e^-$

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note "The Z boson" and ref. LEP-SLC 06). For the Z peak, we report the pole asymmetry defined by  $(3/4)A_e^2$  as determined by the nine-parameter fit to cross-section and lepton forward-backward asymmetry data.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b>1.45 ± 0.25 OUR FIT</b>				
0.89 ± 0.44	1.57	91.2	<sup>1</sup> ABBIENDI	01A OPAL
1.71 ± 0.49	1.57	91.2	ABREU	00F DLPH
1.06 ± 0.58	1.57	91.2	ACCIARRI	00C L3
1.88 ± 0.34	1.57	91.2	<sup>2</sup> BARATE	00C ALEP

<sup>1</sup> ABBIENDI 01A error includes approximately 0.38 due to statistics, 0.16 due to event selection systematics, and 0.18 due to the theoretical uncertainty in  $t$ -channel prediction.

<sup>2</sup> BARATE 00C error includes approximately 0.31 due to statistics, 0.06 due to experimental systematics, and 0.13 due to the theoretical uncertainty in  $t$ -channel prediction.

### $A_{\text{FB}}^{(0,\mu)}$ CHARGE ASYMMETRY IN $e^+ e^- \rightarrow \mu^+ \mu^-$

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note "The Z boson" and ref. LEP-SLC 06). For the Z peak, we report the pole asymmetry defined by  $(3/4)A_e A_\mu$  as determined by the nine-parameter fit to cross-section and lepton forward-backward asymmetry data.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b>1.69 ± 0.13 OUR FIT</b>				
1.59 ± 0.23	1.57	91.2	<sup>1</sup> ABBIENDI	01A OPAL
1.65 ± 0.25	1.57	91.2	ABREU	00F DLPH
1.88 ± 0.33	1.57	91.2	ACCIARRI	00C L3
1.71 ± 0.24	1.57	91.2	<sup>2</sup> BARATE	00C ALEP

• • • We do not use the following data for averages, fits, limits, etc. • • •

9 ± 30	–1.3	20	<sup>3</sup> ABREU	95M DLPH
7 ± 26	–8.3	40	<sup>3</sup> ABREU	95M DLPH
–11 ± 33	–24.1	57	<sup>3</sup> ABREU	95M DLPH
–62 ± 17	–44.6	69	<sup>3</sup> ABREU	95M DLPH
–56 ± 10	–63.5	79	<sup>3</sup> ABREU	95M DLPH
–13 ± 5	–34.4	87.5	<sup>3</sup> ABREU	95M DLPH
–29.0 ± 5.0 – 4.8	–32.1	56.9	<sup>4</sup> ABE	90I VNS
– 9.9 ± 1.5 ± 0.5	–9.2	35	HEGNER	90 JADE
0.05 ± 0.22	0.026	91.14	<sup>5</sup> ABRAMS	89D MRK2
–43.4 ± 17.0	–24.9	52.0	<sup>6</sup> BACALA	89 AMY
–11.0 ± 16.5	–29.4	55.0	<sup>6</sup> BACALA	89 AMY
–30.0 ± 12.4	–31.2	56.0	<sup>6</sup> BACALA	89 AMY
–46.2 ± 14.9	–33.0	57.0	<sup>6</sup> BACALA	89 AMY
–29 ± 13	–25.9	53.3	ADACHI	88C TOPZ
+ 5.3 ± 5.0 ± 0.5	–1.2	14.0	ADEVA	88 MRKJ
–10.4 ± 1.3 ± 0.5	–8.6	34.8	ADEVA	88 MRKJ
–12.3 ± 5.3 ± 0.5	–10.7	38.3	ADEVA	88 MRKJ
–15.6 ± 3.0 ± 0.5	–14.9	43.8	ADEVA	88 MRKJ
– 1.0 ± 6.0	–1.2	13.9	BRAUNSCH...	88D TASS
– 9.1 ± 2.3 ± 0.5	–8.6	34.5	BRAUNSCH...	88D TASS
–10.6 ± 2.2 ± 0.5 – 2.3	–8.9	35.0	BRAUNSCH...	88D TASS
–17.6 ± 4.4 ± 0.5 – 4.3	–15.2	43.6	BRAUNSCH...	88D TASS
– 4.8 ± 6.5 ± 1.0	–11.5	39	BEHREND	87C CELL
–18.8 ± 4.5 ± 1.0	–15.5	44	BEHREND	87C CELL
+ 2.7 ± 4.9	–1.2	13.9	BARTEL	86C JADE
–11.1 ± 1.8 ± 1.0	–8.6	34.4	BARTEL	86C JADE
–17.3 ± 4.8 ± 1.0	–13.7	41.5	BARTEL	86C JADE
–22.8 ± 5.1 ± 1.0	–16.6	44.8	BARTEL	86C JADE
– 6.3 ± 0.8 ± 0.2	–6.3	29	ASH	85 MAC
– 4.9 ± 1.5 ± 0.5	–5.9	29	DERRICK	85 HRS
– 7.1 ± 1.7	–5.7	29	LEVI	83 MRK2
–16.1 ± 3.2	–9.2	34.2	BRANDELIK	82C TASS

<sup>1</sup> ABBIENDI 01A error is almost entirely on account of statistics.

<sup>2</sup> BARATE 00C error is almost entirely on account of statistics.

<sup>3</sup> ABREU 95M perform this measurement using radiative muon-pair events associated with high-energy isolated photons.

<sup>4</sup> ABE 90I measurements in the range  $50 \leq \sqrt{s} \leq 60.8$  GeV.

<sup>5</sup> ABRAMS 89D asymmetry includes both  $9 \mu^+ \mu^-$  and  $15 \tau^+ \tau^-$  events.

<sup>6</sup> BACALA 89 systematic error is about 5%.

### $A_{\text{FB}}^{(0,\tau)}$ CHARGE ASYMMETRY IN $e^+ e^- \rightarrow \tau^+ \tau^-$

OUR FIT is obtained using the fit procedure and correlations as determined by the LEP Electroweak Working Group (see the note "The Z boson" and ref. LEP-SLC 06). For the Z peak, we report the pole asymmetry defined by  $(3/4)A_e A_\tau$  as determined by the nine-parameter fit to cross-section and lepton forward-backward asymmetry data.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b>1.88 ± 0.17 OUR FIT</b>				
1.45 ± 0.30	1.57	91.2	<sup>1</sup> ABBIENDI	01A OPAL
2.41 ± 0.37	1.57	91.2	ABREU	00F DLPH
2.60 ± 0.47	1.57	91.2	ACCIARRI	00C L3
1.70 ± 0.28	1.57	91.2	<sup>2</sup> BARATE	00C ALEP

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-32.8 \pm 6.4 \pm 1.5$	$-32.1$	56.9	<sup>3</sup> ABE	90i	VNS
$-8.1 \pm 2.0 \pm 0.6$	$-9.2$	35	HEGNER	90	JADE
$-18.4 \pm 19.2$	$-24.9$	52.0	<sup>4</sup> BACALA	89	AMY
$-17.7 \pm 26.1$	$-29.4$	55.0	<sup>4</sup> BACALA	89	AMY
$-45.9 \pm 16.6$	$-31.2$	56.0	<sup>4</sup> BACALA	89	AMY
$-49.5 \pm 18.0$	$-33.0$	57.0	<sup>4</sup> BACALA	89	AMY
$-20 \pm 14$	$-25.9$	53.3	ADACHI	88c	TOPZ
$-10.6 \pm 3.1 \pm 1.5$	$-8.5$	34.7	ADEVA	88	MRKJ
$-8.5 \pm 6.6 \pm 1.5$	$-15.4$	43.8	ADEVA	88	MRKJ
$-6.0 \pm 2.5 \pm 1.0$	8.8	34.6	BARTEL	85f	JADE
$-11.8 \pm 4.6 \pm 1.0$	14.8	43.0	BARTEL	85f	JADE
$-5.5 \pm 1.2 \pm 0.5$	$-0.063$	29.0	FERNANDEZ	85	MAC
$-4.2 \pm 2.0$	0.057	29	LEVI	83	MRK2
$-10.3 \pm 5.2$	$-9.2$	34.2	BEHREND	82	CELL
$-0.4 \pm 6.6$	$-9.1$	34.2	BRANDELIK	82c	TASS

<sup>1</sup> ABBIENDI 01a error includes approximately 0.26 due to statistics and 0.14 due to event selection systematics.

<sup>2</sup> BARATE 00c error includes approximately 0.26 due to statistics and 0.11 due to experimental systematics.

<sup>3</sup> ABE 90i measurements in the range  $50 \leq \sqrt{s} \leq 60.8$  GeV.

<sup>4</sup> BACALA 89 systematic error is about 5%.

### $A_{FB}^{(0,\ell)}$ CHARGE ASYMMETRY IN $e^+e^- \rightarrow \ell^+\ell^-$

For the Z peak, we report the pole asymmetry defined by  $(3/4)A_{FB}^2$  as determined by the five-parameter fit to cross-section and lepton forward-backward asymmetry data assuming lepton universality. For details see the note "The Z boson" and ref. LEP-SLC 06.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b><math>1.71 \pm 0.10</math> OUR FIT</b>				
$1.45 \pm 0.17$	1.57	91.2	<sup>1</sup> ABBIENDI 01a	OPAL
$1.87 \pm 0.19$	1.57	91.2	ABREU 00f	DLPH
$1.92 \pm 0.24$	1.57	91.2	ACCIARRI 00c	L3
$1.73 \pm 0.16$	1.57	91.2	<sup>2</sup> BARATE 00c	ALEP

<sup>1</sup> ABBIENDI 01a error includes approximately 0.15 due to statistics, 0.06 due to event selection systematics, and 0.03 due to the theoretical uncertainty in  $t$ -channel prediction.

<sup>2</sup> BARATE 00c error includes approximately 0.15 due to statistics, 0.04 due to experimental systematics, and 0.02 due to the theoretical uncertainty in  $t$ -channel prediction.

### $A_{FB}^{(0,u)}$ CHARGE ASYMMETRY IN $e^+e^- \rightarrow u\bar{u}$

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b><math>4.0 \pm 6.7 \pm 2.8</math></b>	<b>7.2</b>	<b>91.2</b>	<sup>1</sup> ACKERSTAFF 97t	OPAL

<sup>1</sup> ACKERSTAFF 97t measure the forward-backward asymmetry of various fast hadrons made of light quarks. Then using SU(2) isospin symmetry and flavor independence for down and strange quarks authors solve for the different quark types.

### $A_{FB}^{(0,s)}$ CHARGE ASYMMETRY IN $e^+e^- \rightarrow s\bar{s}$

The  $s$ -quark asymmetry is derived from measurements of the forward-backward asymmetry of fast hadrons containing an  $s$  quark.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b><math>9.8 \pm 1.1</math> OUR AVERAGE</b>				
$10.08 \pm 1.13 \pm 0.40$	10.1	91.2	<sup>1</sup> ABREU 00b	DLPH
$6.8 \pm 3.5 \pm 1.1$	10.1	91.2	<sup>2</sup> ACKERSTAFF 97t	OPAL

<sup>1</sup> ABREU 00b tag the presence of an  $s$  quark requiring a high-momentum-identified charged kaon. The  $s$ -quark pole asymmetry is extracted from the charged-kaon asymmetry taking the expected  $d$ - and  $u$ -quark asymmetries from the Standard Model and using the measured values for the  $c$ - and  $b$ -quark asymmetries.

<sup>2</sup> ACKERSTAFF 97t measure the forward-backward asymmetry of various fast hadrons made of light quarks. Then using SU(2) isospin symmetry and flavor independence for down and strange quarks authors solve for the different quark types. The value reported here corresponds then to the forward-backward asymmetry for "down-type" quarks.

### $A_{FB}^{(0,c)}$ CHARGE ASYMMETRY IN $e^+e^- \rightarrow c\bar{c}$

OUR FIT, which is obtained by a simultaneous fit to several  $c$ - and  $b$ -quark measurements as explained in the note "The Z boson" and ref. LEP-SLC 06, refers to the **Z pole** asymmetry. The experimental values, on the other hand, correspond to the measurements carried out at the respective energies.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b><math>7.07 \pm 0.35</math> OUR FIT</b>				
$6.31 \pm 0.93 \pm 0.65$	6.35	91.26	<sup>1</sup> ABDALLAH 04f	DLPH
$5.68 \pm 0.54 \pm 0.39$	6.3	91.25	<sup>2</sup> ABBIENDI 03p	OPAL
$6.45 \pm 0.57 \pm 0.37$	6.10	91.21	<sup>3</sup> HEISTER 02h	ALEP
$6.59 \pm 0.94 \pm 0.35$	6.2	91.235	<sup>4</sup> ABREU 99y	DLPH
$6.3 \pm 0.9 \pm 0.3$	6.1	91.22	<sup>5</sup> BARATE 98o	ALEP
$6.3 \pm 1.2 \pm 0.6$	6.1	91.22	<sup>6</sup> ALEXANDER 97c	OPAL
$8.3 \pm 3.8 \pm 2.7$	6.2	91.24	<sup>7</sup> ADRIANI 92d	L3

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.1 \pm 3.5 \pm 0.5$	$-3.5$	89.43	<sup>1</sup> ABDALLAH 04f	DLPH
$11.0 \pm 2.8 \pm 0.7$	12.3	92.99	<sup>1</sup> ABDALLAH 04f	DLPH
$-6.8 \pm 2.5 \pm 0.9$	$-3.0$	89.51	<sup>2</sup> ABBIENDI 03p	OPAL
$14.6 \pm 2.0 \pm 0.8$	12.2	92.95	<sup>2</sup> ABBIENDI 03p	OPAL
$-12.4 \pm 15.9 \pm 2.0$	$-9.6$	88.38	<sup>3</sup> HEISTER 02h	ALEP
$-2.3 \pm 2.6 \pm 0.2$	$-3.8$	89.38	<sup>3</sup> HEISTER 02h	ALEP
$-0.3 \pm 8.3 \pm 0.6$	0.9	90.21	<sup>3</sup> HEISTER 02h	ALEP
$10.6 \pm 7.7 \pm 0.7$	9.6	92.05	<sup>3</sup> HEISTER 02h	ALEP
$11.9 \pm 2.1 \pm 0.6$	12.2	92.94	<sup>3</sup> HEISTER 02h	ALEP
$12.1 \pm 11.0 \pm 1.0$	14.2	93.90	<sup>3</sup> HEISTER 02h	ALEP
$-4.96 \pm 3.68 \pm 0.53$	$-3.5$	89.434	<sup>4</sup> ABREU 99y	DLPH
$11.80 \pm 3.18 \pm 0.62$	12.3	92.990	<sup>4</sup> ABREU 99y	DLPH
$-1.0 \pm 4.3 \pm 1.0$	$-3.9$	89.37	<sup>5</sup> BARATE 98o	ALEP
$11.0 \pm 3.3 \pm 0.8$	12.3	92.96	<sup>5</sup> BARATE 98o	ALEP
$3.9 \pm 5.1 \pm 0.9$	$-3.4$	89.45	<sup>6</sup> ALEXANDER 97c	OPAL
$15.8 \pm 4.1 \pm 1.1$	12.4	93.00	<sup>6</sup> ALEXANDER 97c	OPAL
$-12.9 \pm 7.8 \pm 5.5$	$-13.6$	35	BEHREND 90d	CELL
$7.7 \pm 13.4 \pm 5.0$	$-22.1$	43	BEHREND 90d	CELL
$-12.8 \pm 4.4 \pm 4.1$	$-13.6$	35	ELSEN 90	JADE
$-10.9 \pm 12.9 \pm 4.6$	$-23.2$	44	ELSEN 90	JADE
$-14.9 \pm 6.7$	$-13.3$	35	OULD-SAADA 89	JADE

<sup>1</sup> ABDALLAH 04f tag  $b$ - and  $c$ -quarks using semileptonic decays combined with charge flow information from the hemisphere opposite to the lepton. Enriched samples of  $c\bar{c}$  and  $b\bar{b}$  events are obtained using lifetime information.

<sup>2</sup> ABBIENDI 03p tag heavy flavors using events with one or two identified leptons. This allows the simultaneous fitting of the  $b$  and  $c$  quark forward-backward asymmetries as well as the average  $B^0\text{-}\bar{B}^0$  mixing.

<sup>3</sup> HEISTER 02h measure simultaneously  $b$  and  $c$  quark forward-backward asymmetries using their semileptonic decays to tag the quark charge. The flavor separation is obtained with a discriminating multivariate analysis.

<sup>4</sup> ABREU 99y tag  $Z \rightarrow b\bar{b}$  and  $Z \rightarrow c\bar{c}$  events by an exclusive reconstruction of several  $D$  meson decay modes ( $D^{*+}$ ,  $D^0$ , and  $D^+$  with their charge-conjugate states).

<sup>5</sup> BARATE 98o tag  $Z \rightarrow c\bar{c}$  events requiring the presence of high-momentum reconstructed  $D^{*+}$ ,  $D^+$ , or  $D^0$  mesons.

<sup>6</sup> ALEXANDER 97c identify the  $b$  and  $c$  events using a  $D/D^*$  tag.

<sup>7</sup> ADRIANI 92d use both electron and muon semileptonic decays.

### $A_{FB}^{(0,b)}$ CHARGE ASYMMETRY IN $e^+e^- \rightarrow b\bar{b}$

OUR FIT, which is obtained by a simultaneous fit to several  $c$ - and  $b$ -quark measurements as explained in the note "The Z boson" and ref. LEP-SLC 06, refers to the **Z pole** asymmetry. The experimental values, on the other hand, correspond to the measurements carried out at the respective energies.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
<b><math>9.92 \pm 0.16</math> OUR FIT</b>				
$9.58 \pm 0.32 \pm 0.14$	9.68	91.231	<sup>1</sup> ABDALLAH 05	DLPH
$10.04 \pm 0.56 \pm 0.25$	9.69	91.26	<sup>2</sup> ABDALLAH 04f	DLPH
$9.72 \pm 0.42 \pm 0.15$	9.67	91.25	<sup>3</sup> ABBIENDI 03p	OPAL
$9.77 \pm 0.36 \pm 0.18$	9.69	91.26	<sup>4</sup> ABBIENDI 02i	OPAL
$9.52 \pm 0.41 \pm 0.17$	9.59	91.21	<sup>5</sup> HEISTER 02h	ALEP
$10.00 \pm 0.27 \pm 0.11$	9.63	91.232	<sup>6</sup> HEISTER 01d	ALEP
$7.62 \pm 1.94 \pm 0.85$	9.64	91.235	<sup>7</sup> ABREU 99y	DLPH
$9.60 \pm 0.66 \pm 0.33$	9.69	91.26	<sup>8</sup> ACCIARRI 99d	L3
$9.31 \pm 1.01 \pm 0.55$	9.65	91.24	<sup>9</sup> ACCIARRI 98u	L3
$9.4 \pm 2.7 \pm 2.2$	9.61	91.22	<sup>10</sup> ALEXANDER 97c	OPAL

• • • We do not use the following data for averages, fits, limits, etc. • • •

$6.37 \pm 1.43 \pm 0.17$	5.8	89.449	<sup>1</sup> ABDALLAH 05	DLPH
$10.41 \pm 1.15 \pm 0.24$	12.1	92.990	<sup>1</sup> ABDALLAH 05	DLPH
$6.7 \pm 2.2 \pm 0.2$	5.7	89.43	<sup>2</sup> ABDALLAH 04f	DLPH
$11.2 \pm 1.8 \pm 0.2$	12.1	92.99	<sup>2</sup> ABDALLAH 04f	DLPH
$4.7 \pm 1.8 \pm 0.1$	5.9	89.51	<sup>3</sup> ABBIENDI 03p	OPAL
$10.3 \pm 1.5 \pm 0.2$	12.0	92.95	<sup>3</sup> ABBIENDI 03p	OPAL
$5.82 \pm 1.53 \pm 0.12$	5.9	89.50	<sup>4</sup> ABBIENDI 02i	OPAL
$12.21 \pm 1.23 \pm 0.25$	12.0	92.91	<sup>4</sup> ABBIENDI 02i	OPAL
$-13.1 \pm 13.5 \pm 1.0$	3.2	88.38	<sup>5</sup> HEISTER 02h	ALEP
$5.5 \pm 1.9 \pm 0.1$	5.6	89.38	<sup>5</sup> HEISTER 02h	ALEP
$-0.4 \pm 6.7 \pm 0.8$	7.5	90.21	<sup>5</sup> HEISTER 02h	ALEP
$11.1 \pm 6.4 \pm 0.5$	11.0	92.05	<sup>5</sup> HEISTER 02h	ALEP
$10.4 \pm 1.5 \pm 0.3$	12.0	92.94	<sup>5</sup> HEISTER 02h	ALEP
$13.8 \pm 9.3 \pm 1.1$	12.9	93.90	<sup>5</sup> HEISTER 02h	ALEP
$4.36 \pm 1.19 \pm 0.11$	5.8	89.472	<sup>6</sup> HEISTER 01d	ALEP
$11.72 \pm 0.97 \pm 0.11$	12.0	92.950	<sup>6</sup> HEISTER 01d	ALEP
$5.67 \pm 7.56 \pm 1.17$	5.7	89.434	<sup>7</sup> ABREU 99y	DLPH
$8.82 \pm 6.33 \pm 1.22$	12.1	92.990	<sup>7</sup> ABREU 99y	DLPH
$6.11 \pm 2.93 \pm 0.43$	5.9	89.50	<sup>8</sup> ACCIARRI 99d	L3
$13.71 \pm 2.40 \pm 0.44$	12.2	93.10	<sup>8</sup> ACCIARRI 99d	L3
$4.95 \pm 5.23 \pm 0.40$	5.8	89.45	<sup>9</sup> ACCIARRI 98u	L3
$11.37 \pm 3.99 \pm 0.65$	12.1	92.99	<sup>9</sup> ACCIARRI 98u	L3
$-8.6 \pm 10.8 \pm 2.9$	5.8	89.45	<sup>10</sup> ALEXANDER 97c	OPAL
$-2.1 \pm 9.0 \pm 2.6$	12.1	93.00	<sup>10</sup> ALEXANDER 97c	OPAL
$-71 \pm 34 \pm 7$	$-58$	58.3	SHIMONAKA 91	TOPZ
$-22.2 \pm 7.7 \pm 3.5$	$-26.0$	35	BEHREND 90d	CELL

# Gauge & Higgs Boson Particle Listings

## Z

-49.1 ± 16.0 ± 5.0	-39.7	43	BEHREND	90D	CELL
-28 ± 11	-23	35	BRAUNSCH...	90	TASS
-16.6 ± 7.7 ± 4.8	-24.3	35	ELSEN	90	JADE
-33.6 ± 22.2 ± 5.2	-39.9	44	ELSEN	90	JADE
3.4 ± 7.0 ± 3.5	-16.0	29.0	BAND	89	MAC
-72 ± 28 ± 13	-56	55.2	SAGAWA	89	AMY

<sup>1</sup> ABDALLAH 05 obtain an enriched samples of  $b\bar{b}$  events using lifetime information. The quark (or antiquark) charge is determined with a neural network using the secondary vertex charge, the jet charge and particle identification.

<sup>2</sup> ABDALLAH 04F tag  $b$ - and  $c$ -quarks using semileptonic decays combined with charge flow information from the hemisphere opposite to the lepton. Enriched samples of  $c\bar{c}$  and  $b\bar{b}$  events are obtained using lifetime information.

<sup>3</sup> ABBIENDI 03P tag heavy flavors using events with one or two identified leptons. This allows the simultaneous fitting of the  $b$  and  $c$  quark forward-backward asymmetries as well as the average  $B^0\text{-}\bar{B}^0$  mixing.

<sup>4</sup> ABBIENDI 02i tag  $Z^0 \rightarrow b\bar{b}$  decays using a combination of secondary vertex and lepton tags. The sign of the  $b$ -quark charge is determined using an inclusive tag based on jet, vertex, and kaon charges.

<sup>5</sup> HEISTER 02H measure simultaneously  $b$  and  $c$  quark forward-backward asymmetries using their semileptonic decays to tag the quark charge. The flavor separation is obtained with a discriminating multivariate analysis.

<sup>6</sup> HEISTER 01D tag  $Z \rightarrow b\bar{b}$  events using the impact parameters of charged tracks complemented with information from displaced vertices, event shape variables, and lepton identification. The  $b$ -quark direction and charge is determined using the hemisphere charge method along with information from fast kaon tagging and charge estimators of primary and secondary vertices. The change in the quoted value due to variation of  $A_{FB}^e$  and  $R_b$  is given as  $+0.103 (A_{FB}^e - 0.0651) - 0.440 (R_b - 0.21585)$ .

<sup>7</sup> ABREU 99Y tag  $Z \rightarrow b\bar{b}$  and  $Z \rightarrow c\bar{c}$  events by an exclusive reconstruction of several  $D$  meson decay modes ( $D^{*+}$ ,  $D^0$ , and  $D^+$  with their charge-conjugate states).

<sup>8</sup> ACCIARRI 99D tag  $Z \rightarrow b\bar{b}$  events using high  $p$  and  $p_T$  leptons. The analysis determines simultaneously a mixing parameter  $\chi_b = 0.1192 \pm 0.0068 \pm 0.0051$  which is used to correct the observed asymmetry.

<sup>9</sup> ACCIARRI 98U tag  $Z \rightarrow b\bar{b}$  events using lifetime and measure the jet charge using the hemisphere charge.

<sup>10</sup> ALEXANDER 97C identify the  $b$  and  $c$  events using a  $D/D^*$  tag.

### CHARGE ASYMMETRY IN $e^+e^- \rightarrow q\bar{q}$

Summed over five lighter flavors.

Experimental and Standard Model values are somewhat event-selection dependent. Standard Model expectations contain some assumptions on  $B^0\text{-}\bar{B}^0$  mixing and on other electroweak parameters.

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.76 ± 0.12 ± 0.15		91.2	<sup>1</sup> ABREU 92i	DLPH
4.0 ± 0.4 ± 0.63	4.0	91.3	<sup>2</sup> ACTON 92L	OPAL
9.1 ± 1.4 ± 1.6	9.0	57.9	ADACHI 91	TOPZ
-0.84 ± 0.15 ± 0.04		91	DECAMP 91B	ALEP
8.3 ± 2.9 ± 1.9	8.7	56.6	STUART 90	AMY
11.4 ± 2.2 ± 2.1	8.7	57.6	ABE 89L	VNS
6.0 ± 1.3	5.0	34.8	GREENSHAW 89	JADE
8.2 ± 2.9	8.5	43.6	GREENSHAW 89	JADE

<sup>1</sup> ABREU 92i has 0.14 systematic error due to uncertainty of quark fragmentation.

<sup>2</sup> ACTON 92L use the weight function method on 259k selected  $Z \rightarrow$  hadrons events. The systematic error includes a contribution of 0.2 due to  $B^0\text{-}\bar{B}^0$  mixing effect, 0.4 due to Monte Carlo (MC) fragmentation uncertainties and 0.3 due to MC statistics. ACTON 92L derive a value of  $\sin^2 \theta_{\text{eff}}^W$  to be  $0.2321 \pm 0.0017 \pm 0.0028$ .

### CHARGE ASYMMETRY IN $p\bar{p} \rightarrow Z \rightarrow e^+e^-$

ASYMMETRY (%)	STD. MODEL	$\sqrt{s}$ (GeV)	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.2 ± 5.9 ± 0.4		91	ABE	91E CDF

### ANOMALOUS $ZZ\gamma$ , $Z\gamma\gamma$ , AND $ZZV$ COUPLINGS

### ANOMALOUS $ZZ\gamma$ , $Z\gamma\gamma$ , AND $ZZV$ COUPLINGS

Revised September 2013 by M.W. Grünewald (U. College Dublin and U. Ghent) and A. Gurtu (Formerly Tata Inst.).

In on-shell  $Z\gamma$  production, deviations from the Standard Model for the  $Z\gamma\gamma^*$  and  $Z\gamma Z^*$  couplings may be described in terms of eight parameters,  $h_i^V$  ( $i = 1, 4$ ;  $V = \gamma, Z$ ) [1]. The

parameters  $h_i^\gamma$  describe the  $Z\gamma\gamma^*$  couplings and the parameters  $h_i^Z$  the  $Z\gamma Z^*$  couplings. In this formalism  $h_1^V$  and  $h_2^V$  lead to  $CP$ -violating and  $h_3^V$  and  $h_4^V$  to  $CP$ -conserving effects. All these anomalous contributions to the cross section increase rapidly with center-of-mass energy. In order to ensure unitarity, these parameters are usually described by a form-factor representation,  $h_i^V(s) = h_{i0}^V/(1 + s/\Lambda^2)^n$ , where  $\Lambda$  is the energy scale for the manifestation of a new phenomenon and  $n$  is a sufficiently large power. By convention one uses  $n = 3$  for  $h_{1,3}^V$  and  $n = 4$  for  $h_{2,4}^V$ . Usually limits on  $h_i^V$ 's are put assuming some value of  $\Lambda$ , sometimes  $\infty$ .

In on-shell  $ZZ$  production, deviations from the Standard Model for the  $ZZ\gamma^*$  and  $ZZZ^*$  couplings may be described by means of four anomalous couplings  $f_i^V$  ( $i = 4, 5$ ;  $V = \gamma, Z$ ) [2]. As above, the parameters  $f_i^\gamma$  describe the  $ZZ\gamma^*$  couplings and the parameters  $f_i^Z$  the  $ZZZ^*$  couplings. The anomalous couplings  $f_5^V$  lead to violation of  $C$  and  $P$  symmetries while  $f_4^V$  introduces  $CP$  violation. Also here, formfactors depending on a scale  $\Lambda$  are used.

All these couplings  $h_i^V$  and  $f_i^V$  are zero at tree level in the Standard Model; they are measured in  $e^+e^-$ ,  $p\bar{p}$  and  $pp$  collisions at LEP, Tevatron and LHC.

### References

1. U. Baur and E.L. Berger Phys. Rev. **D47**, 4889 (1993).
2. K. Hagiwara *et al.*, Nucl. Phys. **B282**, 253 (1987).

### $h_i^V$

Combining the LEP-2 results taking into account the correlations, the following 95% CL limits are derived [SCHAEL 13A]:

$$\begin{aligned} -0.12 < h_1^Z < +0.11, & \quad -0.07 < h_2^Z < +0.07, \\ -0.19 < h_3^Z < +0.06, & \quad -0.04 < h_4^Z < +0.13, \\ -0.05 < h_1^\gamma < +0.05, & \quad -0.04 < h_2^\gamma < +0.02, \\ -0.05 < h_3^\gamma < +0.00, & \quad +0.01 < h_4^\gamma < +0.05. \end{aligned}$$

Some of the recent results from the Tevatron and LHC experiments individually surpass the combined LEP-2 results in precision (see below).

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<sup>1</sup> KHACHATRYAN 15AC	CMS	$E_{\text{cm}}^{pp} = 8$ TeV	
<sup>2</sup> CHATRCHYAN 14AB	CMS	$E_{\text{cm}}^{pp} = 7$ TeV	
<sup>3</sup> AAD	13AN ATLS	$E_{\text{cm}}^{pp} = 7$ TeV	
<sup>4</sup> CHATRCHYAN 13BI	CMS	$E_{\text{cm}}^{pp} = 7$ TeV	
<sup>5</sup> ABAZOV	12s D0	$E_{\text{cm}}^{pp} = 1.96$ TeV	
<sup>6</sup> AALTONEN	11s CDF	$E_{\text{cm}}^{pp} = 1.96$ TeV	
<sup>7</sup> CHATRCHYAN 11M	CMS	$E_{\text{cm}}^{pp} = 7$ TeV	
<sup>8</sup> ABAZOV	09L D0	$E_{\text{cm}}^{pp} = 1.96$ TeV	
<sup>9</sup> ABAZOV	07M D0	$E_{\text{cm}}^{pp} = 1.96$ TeV	
<sup>10</sup> ABDALLAH	07C DLPH	$E_{\text{cm}}^{ee} = 183\text{--}208$ GeV	
<sup>11</sup> ACHARD	04H L3	$E_{\text{cm}}^{ee} = 183\text{--}208$ GeV	
<sup>12</sup> ABBIENDI, G	00C OPAL	$E_{\text{cm}}^{ee} = 189$ GeV	
<sup>13</sup> ABBOTT	98M D0	$E_{\text{cm}}^{pp} = 1.8$ TeV	
<sup>14</sup> ABREU	98K DLPH	$E_{\text{cm}}^{ee} = 161, 172$ GeV	

<sup>1</sup> KHACHATRYAN 15AC study  $Z\gamma$  events in 8 TeV  $pp$  interactions, where the  $Z$  decays into 2 same-flavor, opposite sign leptons ( $e$  or  $\mu$ ) and a photon with  $p_T > 15$  GeV. The  $p_T$  of a lepton is required to be  $> 20$  GeV/ $c$ , their effective mass  $> 50$  GeV, and the photon should have a separation  $\Delta R > 0.7$  with each lepton. The observed  $p_T$  distribution of the photons is used to extract the 95% C.L. limits:  $-3.8 \times 10^{-3} < h_3^Z < 3.7 \times 10^{-3}$ ,  $-3.1 \times 10^{-5} < h_4^Z < 3.0 \times 10^{-5}$ ,  $-4.6 \times 10^{-3} < h_3^\gamma < 4.6 \times 10^{-3}$ ,  $-3.6 \times 10^{-5} < h_4^\gamma < 3.5 \times 10^{-5}$ .

- <sup>2</sup> CHATRCHYAN 14AB measure  $Z\gamma$  production cross section for  $p_T^{\gamma} > 15$  GeV and  $R(\ell\gamma) > 0.7$ , which is the separation between the  $\gamma$  and the final state charged lepton ( $e$  or  $\mu$ ) in the azimuthal angle-pseudorapidity ( $\phi - \eta$ ) plane. The di-lepton mass is required to be  $> 50$  GeV. After background subtraction the number of  $e\bar{e}\gamma$  and  $\mu\bar{\mu}\gamma$  events is determined to be  $3160 \pm 120$  and  $5030 \pm 233$  respectively, compatible with expectations from the SM. This leads to a 95% CL limits of  $-1 \times 10^{-2} < h_3^Z < 1 \times 10^{-2}$ ,  $-9 \times 10^{-5} < h_4^Z < 9 \times 10^{-5}$ ,  $-9 \times 10^{-3} < h_5^Z < 9 \times 10^{-3}$ ,  $-8 \times 10^{-5} < h_4^Z < 8 \times 10^{-5}$ , assuming  $h_1^V$  and  $h_2^V$  have SM values,  $V = \gamma$  or  $Z$ .
- <sup>3</sup> AAD 13AN study  $Z\gamma$  production in  $pp$  collisions. In events with no additional jet, 1417 (2031)  $Z$  decays to electron (muon) pairs are selected, with an expected background of  $156 \pm 54$  ( $244 \pm 64$ ) events, as well as 662  $Z$  decays to neutrino pairs with an expected background of  $302 \pm 42$  events. Analysing the photon  $p_T$  spectrum above 100 GeV yields the 95% C.L. limits:  $-0.013 < h_3^Z < 0.014$ ,  $-8.7 \times 10^{-5} < h_4^Z < 8.7 \times 10^{-5}$ ,  $-0.015 < h_3^Z < 0.016$ ,  $-9.4 \times 10^{-5} < h_4^Z < 9.2 \times 10^{-5}$ . Supersedes AAD 12BX.
- <sup>4</sup> CHATRCHYAN 13BI determine the  $Z\gamma \rightarrow \nu\bar{\nu}\gamma$  cross section by selecting events with a photon of  $E_T > 145$  GeV and a  $p_T > 130$  GeV. 73 candidate events are observed with an expected SM background of  $30.2 \pm 6.5$ . The  $E_T$  spectrum of the photon is used to set 95% C.L. limits as follows:  $|h_3^Z| < 2.7 \times 10^{-3}$ ,  $|h_4^Z| < 1.3 \times 10^{-5}$ ,  $|h_3^Z| < 2.9 \times 10^{-3}$ ,  $|h_4^Z| < 1.5 \times 10^{-5}$ .
- <sup>5</sup> ABAZOV 12S study  $Z\gamma$  production in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV using  $6.2 \text{ fb}^{-1}$  of data where the  $Z$  decays to electron (muon) pairs and the photon has at least 10 GeV of transverse momentum. In data, 304 (308) di-electron (di-muon) events are observed with an expected background of  $255 \pm 16$  ( $285 \pm 24$ ) events. Based on the photon  $p_T$  spectrum, and including also earlier data and the  $Z \rightarrow \nu\bar{\nu}$  decay mode (from ABAZOV 09L), the following 95% C.L. limits are reported:  $|h_{03}^Z| < 0.026$ ,  $|h_{04}^Z| < 0.0013$ ,  $|h_{03}^Z| < 0.027$ ,  $|h_{04}^Z| < 0.0014$  for a form factor scale of  $\Lambda = 1.5$  TeV.
- <sup>6</sup> AALTONEN 11S study  $Z\gamma$  events in  $p\bar{p}$  interactions at  $\sqrt{s} = 1.96$  TeV with integrated luminosity  $5.1 \text{ fb}^{-1}$  for  $Z \rightarrow e^+e^-/\mu^+\mu^-$  and  $4.9 \text{ fb}^{-1}$  for  $Z \rightarrow \nu\bar{\nu}$ . For the charged lepton case, the two leptons must be of the same flavor with the transverse momentum/energy of one  $> 20$  GeV and the other  $> 10$  GeV. The isolated photon must have  $E_T > 50$  GeV. They observe 91 events with  $87.2 \pm 7.8$  events expected from standard model processes. For the  $\nu\bar{\nu}$  case they require solitary photons with  $E_T > 25$  GeV and missing  $E_T > 25$  GeV and observe 85 events with standard model expectation of  $85.9 \pm 5.6$  events. Taking the form factor  $\Lambda = 1.5$  TeV they derive 95% C.L. limits as  $|h_3^Z| < 0.022$  and  $|h_4^Z| < 0.0009$ .
- <sup>7</sup> CHATRCHYAN 11M study  $Z\gamma$  production in  $pp$  collisions at  $\sqrt{s} = 7$  TeV using 36  $\text{pb}^{-1}$   $pp$  data, where the  $Z$  decays to  $e^+e^-$  or  $\mu^+\mu^-$ . The total cross sections are measured for photon transverse energy  $E_T^{\gamma} > 10$  GeV and spatial separation from charged leptons in the plane of pseudo rapidity and azimuthal angle  $\Delta R(\ell, \gamma) > 0.7$  with the dilepton invariant mass requirement of  $M_{\ell\ell} > 50$  GeV. The number of  $e^+e^-\gamma$  and  $\mu^+\mu^-\gamma$  candidates is 81 and 90 with estimated backgrounds of  $20.5 \pm 2.5$  and  $27.3 \pm 3.2$  events respectively. The 95% CL limits for  $ZZ\gamma$  couplings are  $-0.05 < h_3^Z < 0.06$  and  $-0.0005 < h_4^Z < 0.0005$ , and for  $Z\gamma\gamma$  couplings are  $-0.07 < h_3^Z < 0.07$  and  $-0.0005 < h_4^Z < 0.0006$ .
- <sup>8</sup> ABAZOV 09L study  $Z\gamma, Z \rightarrow \nu\bar{\nu}$  production in  $p\bar{p}$  collisions at 1.96 TeV C.M. energy. They select 51 events with a photon of transverse energy  $E_T$  larger than 90 GeV, with an expected background of 17 events. Based on the photon  $E_T$  spectrum and including also  $Z$  decays to charged leptons (from ABAZOV 07M), the following 95% CL limits are reported:  $|h_{30}^Z| < 0.033$ ,  $|h_{40}^Z| < 0.0017$ ,  $|h_{30}^Z| < 0.033$ ,  $|h_{40}^Z| < 0.0017$ .
- <sup>9</sup> ABAZOV 07M use 968  $p\bar{p} \rightarrow e^+e^-/\mu^+\mu^-\gamma X$  candidates, at 1.96 TeV center of mass energy, to tag  $p\bar{p} \rightarrow Z\gamma$  events by requiring  $E_T(\gamma) > 7$  GeV, lepton-gamma separation  $\Delta R_{\ell\gamma} > 0.7$ , and dilepton invariant mass  $> 30$  GeV. The cross section is in agreement with the SM prediction. Using these  $Z\gamma$  events they obtain 95% C.L. limits on each  $h_i^V$ , keeping all others fixed at their SM values. They report:  $-0.083 < h_{30}^Z < 0.082$ ,  $-0.0053 < h_{40}^Z < 0.0054$ ,  $-0.085 < h_{30}^Z < 0.084$ ,  $-0.0053 < h_{40}^Z < 0.0054$ , for the form factor scale  $\Lambda = 1.2$  TeV.
- <sup>10</sup> Using data collected at  $\sqrt{s} = 183\text{--}208$  GeV, ABDALLAH 07c select 1,877  $e^+e^- \rightarrow Z\gamma$  events with  $Z \rightarrow q\bar{q}$  or  $\nu\bar{\nu}$ , 171  $e^+e^- \rightarrow ZZ$  events with  $Z \rightarrow q\bar{q}$  or lepton pair (except an explicit  $\tau$  pair), and 74  $e^+e^- \rightarrow Z\gamma^*$  events with a  $q\bar{q}\mu^+\mu^-$  or  $q\bar{q}e^+e^-$  signature, to derive 95% CL limits on  $h_i^V$ . Each limit is derived with other parameters set to zero. They report:  $-0.23 < h_1^Z < 0.23$ ,  $-0.30 < h_2^Z < 0.16$ ,  $-0.14 < h_1^Z < 0.14$ ,  $-0.049 < h_3^Z < 0.044$ .
- <sup>11</sup> ACHARD 04H select 3515  $e^+e^- \rightarrow Z\gamma$  events with  $Z \rightarrow q\bar{q}$  or  $\nu\bar{\nu}$  at  $\sqrt{s} = 189\text{--}209$  GeV to derive 95% CL limits on  $h_i^V$ . For deriving each limit the other parameters are fixed at zero. They report:  $-0.153 < h_1^Z < 0.141$ ,  $-0.087 < h_2^Z < 0.079$ ,  $-0.220 < h_3^Z < 0.112$ ,  $-0.068 < h_4^Z < 0.148$ ,  $-0.057 < h_1^Z < 0.057$ ,  $-0.050 < h_2^Z < 0.023$ ,  $-0.059 < h_3^Z < 0.004$ ,  $-0.004 < h_4^Z < 0.042$ .
- <sup>12</sup> ABBIENDI, G 00C study  $e^+e^- \rightarrow Z\gamma$  events (with  $Z \rightarrow q\bar{q}$  and  $Z \rightarrow \nu\bar{\nu}$ ) at 189 GeV to obtain the central values (and 95% CL limits) of these couplings:  $h_1^Z = 0.000 \pm 0.100$  ( $-0.190, 0.190$ ),  $h_2^Z = 0.000 \pm 0.068$  ( $-0.128, 0.128$ ),  $h_3^Z = -0.074 \pm 0.102$  ( $-0.269, 0.119$ ),  $h_4^Z = 0.046 \pm 0.068$  ( $-0.084, 0.175$ ),  $h_1^Z = 0.000 \pm 0.061$  ( $-0.115, 0.115$ ),  $h_2^Z = 0.000 \pm 0.041$  ( $-0.077, 0.077$ ),  $h_3^Z = -0.080 \pm 0.039$  ( $-0.164, -0.006$ ),  $h_4^Z = 0.064 \pm 0.033$  ( $-0.030, +0.134$ ). The results are derived assuming that only one coupling at a time is different from zero.
- <sup>13</sup> ABBOTT 98M study  $p\bar{p} \rightarrow Z\gamma + X$ , with  $Z \rightarrow e^+e^-, \mu^+\mu^-, \nu\bar{\nu}$  at 1.8 TeV, to obtain 95% CL limits at  $\Lambda = 750$  GeV:  $|h_{30}^Z| < 0.36$ ,  $|h_{40}^Z| < 0.05$  (keeping  $h_1^Z = 0$ ), and  $|h_{30}^Z| < 0.37$ ,  $|h_{40}^Z| < 0.05$  (keeping  $h_2^Z = 0$ ). Limits on the  $CP$ -violating couplings are  $|h_{10}^Z| < 0.36$ ,  $|h_{20}^Z| < 0.05$  (keeping  $h_1^Z = 0$ ), and  $|h_{10}^Z| < 0.37$ ,  $|h_{20}^Z| < 0.05$  (keeping  $h_2^Z = 0$ ).
- <sup>14</sup> ABREU 98K determine a 95% CL upper limit on  $\sigma(e^+e^- \rightarrow \gamma + \text{invisible particles}) < 2.5 \text{ pb}$  using 161 and 172 GeV data. This is used to set 95% CL limits on  $|h_{30}^Z| < 0.8$  and  $|h_{30}^Z| < 1.3$ , derived at a scale  $\Lambda = 1$  TeV and with  $n = 3$  in the form factor representation.
- $f_i^V$**   
Combining the LEP-2 results taking into account the correlations, the following 95% CL limits are derived [SCHAE 13A]:
- $$\begin{aligned} -0.28 < f_4^Z < +0.32, & \quad -0.34 < f_5^Z < +0.35, \\ -0.17 < f_4^Z < +0.19, & \quad -0.35 < f_5^Z < +0.32. \end{aligned}$$
- Some of the recent results from the Tevatron and LHC experiments individually surpass the combined LEP-2 results in precision (see below).
- | VALUE   | DOCUMENT ID      | TECN | COMMENT                                    |
|---|------------------|------|--|
| • • • We do not use the following data for averages, fits, limits, etc. • • • |                  |      |  |
| 1   | KHACHATRYAN 15B  | CMS  | $E_{\text{cm}}^{pp} = 8$ TeV               |
| 2   | KHACHATRYAN 15BC | CMS  | $E_{\text{cm}}^{pp} = 7, 8$ TeV            |
| 3   | AAD 13Z          | ATLS | $E_{\text{cm}}^{pp} = 7$ TeV               |
| 4   | CHATRCHYAN 13B   | CMS  | $E_{\text{cm}}^{pp} = 7$ TeV               |
| 5   | SCHAE 09         | ALEP | $E_{\text{cm}}^{ee} = 192\text{--}209$ GeV |
| 6   | ABAZOV 08K       | D0   | $E_{\text{cm}}^{pp} = 1.96$ TeV            |
| 7   | ABDALLAH 07C     | DLPH | $E_{\text{cm}}^{ee} = 183\text{--}208$ GeV |
| 8   | ABBIENDI 04C     | OPAL |  |
| 9   | ACHARD 03D       | L3   |  |
- <sup>1</sup> KHACHATRYAN 15B study  $ZZ$  production in 8 TeV  $pp$  collisions. In the decay modes  $ZZ \rightarrow 4e, 4\mu, 2e2\mu, 54, 75, 148$  events are observed, with an expected background of  $2.2 \pm 0.9, 1.2 \pm 0.6$ , and  $2.4 \pm 1.0$  events, respectively. Analysing the 4-lepton invariant mass spectrum in the range from 110 GeV to 1200 GeV, the following 95% C.L. limits are obtained:  $|f_4^Z| < 0.004$ ,  $|f_5^Z| < 0.004$ ,  $|f_4^Z| < 0.005$ ,  $|f_5^Z| < 0.005$ .
- <sup>2</sup> KHACHATRYAN 15BC use the cross section measurement of the final state  $pp \rightarrow Z \rightarrow 2\ell 2\nu$  ( $\ell$  being an electron or a muon) at 7 and 8 TeV to put limits on these triple gauge couplings. Effective mass of the charged lepton pair is required to be in the range 83.5–98.5 GeV and the dilepton  $p_T > 45$  GeV. The reduced missing  $E_T$  is required to be  $> 65$  GeV, which takes into account the fake missing  $E_T$  due to detector effects. The numbers of  $e^+e^-$  and  $\mu^+\mu^-$  events selected are 35 and 40 at 7 TeV and 176 and 271 at 8 TeV respectively. The production cross sections so obtained are in agreement with SM predictions. The following 95% C.L. limits are set:  $-0.0028 < f_4^Z < 0.0032$ ,  $-0.0037 < f_4^Z < 0.0033$ ,  $-0.0029 < f_5^Z < 0.0031$ ,  $-0.0033 < f_5^Z < 0.0037$ . Combining with previous results (KHACHATRYAN 15B and CHATRCHYAN 13B) which include 7 TeV and 8 TeV data on the final states  $pp \rightarrow ZZ \rightarrow 2\ell 2\ell'$  where  $\ell$  and  $\ell'$  are an electron or a muon, the best limits are  $-0.0022 < f_4^Z < 0.0026$ ,  $-0.0029 < f_4^Z < 0.0026$ ,  $-0.0023 < f_5^Z < 0.0023$ ,  $-0.0026 < f_5^Z < 0.0027$ .
- <sup>3</sup> AAD 13Z study  $ZZ$  production in  $pp$  collisions at  $\sqrt{s} = 7$  TeV. In the  $ZZ \rightarrow \ell^+\ell^-\ell'^+\ell'^-$  final state they observe a total of 66 events with an expected background of  $0.9 \pm 1.3$ . In the  $ZZ \rightarrow \ell^+\ell^-\nu\bar{\nu}$  final state they observe a total of 87 events with an expected background of  $46.9 \pm 5.2$ . The limits on anomalous TGCs are determined using the observed and expected numbers of these  $ZZ$  events binned in  $p_T^Z$ . The 95% C.L. are as follows: for form factor scale  $\Lambda = \infty$ ,  $-0.015 < f_4^Z < 0.015$ ,  $-0.013 < f_4^Z < 0.013$ ,  $-0.016 < f_5^Z < 0.015$ ,  $-0.013 < f_5^Z < 0.013$ ; for form factor scale  $\Lambda = 3$  TeV,  $-0.022 < f_4^Z < 0.023$ ,  $-0.019 < f_4^Z < 0.019$ ,  $-0.023 < f_5^Z < 0.023$ ,  $-0.020 < f_5^Z < 0.019$ .
- <sup>4</sup> CHATRCHYAN 13B study  $ZZ$  production in  $pp$  collisions and select 54  $ZZ$  candidates in the  $Z$  decay channel with electrons or muons with an expected background of  $1.4 \pm 0.5$  events. The resulting 95% C.L. ranges are:  $-0.013 < f_4^Z < 0.015$ ,  $-0.011 < f_4^Z < 0.012$ ,  $-0.014 < f_5^Z < 0.014$ ,  $-0.012 < f_5^Z < 0.012$ .
- <sup>5</sup> Using data collected in the center of mass energy range 192–209 GeV, SCHAE 09 select 318  $e^+e^- \rightarrow ZZ$  events with 319.4 expected from the standard model. Using this data they derive the following 95% CL limits:  $-0.321 < f_4^Z < 0.318$ ,  $-0.534 < f_4^Z < 0.534$ ,  $-0.724 < f_5^Z < 0.733$ ,  $-1.194 < f_5^Z < 1.190$ .
- <sup>6</sup> ABAZOV 08K search for  $ZZ$  and  $Z\gamma^*$  events with  $1 \text{ fb}^{-1}$   $p\bar{p}$  data at  $\sqrt{s} = 1.96$  TeV in  $(e)(e), (\mu\mu)(\mu\mu), (e)(\mu\mu)$  final states requiring the lepton pair masses to be  $> 30$  GeV. They observe 1 event, which is consistent with an expected signal of  $1.71 \pm 0.15$  events and a background of  $0.13 \pm 0.03$  events. From this they derive the following limits, for a form factor ( $\Lambda$ ) value of 1.2 TeV:  $-0.28 < f_{40}^Z < 0.28$ ,  $-0.31 < f_{50}^Z < 0.29$ ,  $-0.26 < f_{40}^Z < 0.26$ ,  $-0.30 < f_{50}^Z < 0.28$ .
- <sup>7</sup> Using data collected at  $\sqrt{s} = 183\text{--}208$  GeV, ABDALLAH 07c select 171  $e^+e^- \rightarrow ZZ$  events with  $Z \rightarrow q\bar{q}$  or lepton pair (except an explicit  $\tau$  pair), and 74  $e^+e^- \rightarrow Z\gamma^*$  events with a  $q\bar{q}\mu^+\mu^-$  or  $q\bar{q}e^+e^-$  signature, to derive 95% CL limits on  $f_i^V$ . Each limit is derived with other parameters set to zero. They report:  $-0.40 < f_4^Z < 0.42$ ,  $-0.38 < f_5^Z < 0.62$ ,  $-0.23 < f_4^Z < 0.25$ ,  $-0.52 < f_5^Z < 0.48$ .
- <sup>8</sup> ABBIENDI 04C study  $ZZ$  production in  $e^+e^-$  collisions in the C.M. energy range 190–209 GeV. They select 340 events with an expected background of 180 events. Including the ABBIENDI 00W data at 183 and 189 GeV (118 events with an expected background of 65 events) they report the following 95% CL limits:  $-0.45 < f_4^Z < 0.58$ ,  $-0.94 < f_5^Z < 0.25$ ,  $-0.32 < f_4^Z < 0.33$ , and  $-0.71 < f_5^Z < 0.59$ .
- <sup>9</sup> ACHARD 03D study  $Z$ -boson pair production in  $e^+e^-$  collisions in the C.M. energy range 200–209 GeV. They select 549 events with an expected background of 432 events. Including the ACCIARRI 99G and ACCIARRI 99O data (183 and 189 GeV respectively, 286

# Gauge & Higgs Boson Particle Listings

## Z

events with an expected background of 241 events) and the 192–202 GeV ACCIARRI 011 results (656 events, expected background of 512 events), they report the following 95% CL limits:  $-0.48 \leq f_4^Z \leq 0.46$ ,  $-0.36 \leq f_5^Z \leq 1.03$ ,  $-0.28 \leq f_4^\gamma \leq 0.28$ , and  $-0.40 \leq f_5^\gamma \leq 0.47$ .

### ANOMALOUS W/Z QUARTIC COUPLINGS

#### ANOMALOUS W/Z QUARTIC COUPLINGS (QGCS)

Revised November 2015 by M.W. Gr  nwald (U. College Dublin) and A. Gurtu (Formerly Tata Inst.).

Quartic couplings,  $WWZZ$ ,  $WWZ\gamma$ ,  $WW\gamma\gamma$ , and  $ZZ\gamma\gamma$ , were studied at LEP and Tevatron at energies at which the Standard Model predicts negligible contributions to multiboson production. Thus, to parametrize limits on these couplings, an effective theory approach is adopted which supplements the Standard Model Lagrangian with higher dimensional operators which include quartic couplings. The LEP collaborations chose the lowers dimensional representation of operators (dimension 6) which presumes the  $SU(2) \times U(1)$  gauge symmetry is broken by means other than the conventional Higgs scalar doublet [1–3]. In this representation possible quartic couplings,  $a_0, a_c, a_n$ , are expressed in terms of the following dimension-6 operators [1,2];

$$\begin{aligned} L_6^0 &= -\frac{e^2}{16\Lambda^2} a_0 F^{\mu\nu} F_{\mu\nu} \vec{W}^\alpha \cdot \vec{W}_\alpha \\ L_6^c &= -\frac{e^2}{16\Lambda^2} a_c F^{\mu\alpha} F_{\mu\beta} \vec{W}^\beta \cdot \vec{W}_\alpha \\ L_6^n &= -i\frac{e^2}{16\Lambda^2} a_n \epsilon_{ijk} W_{\mu\alpha}^{(i)} W_\nu^{(j)} W^{(k)\alpha} F^{\mu\nu} \\ \tilde{L}_6^0 &= -\frac{e^2}{16\Lambda^2} \tilde{a}_0 F^{\mu\nu} \tilde{F}_{\mu\nu} \vec{W}^\alpha \cdot \vec{W}_\alpha \\ \tilde{L}_6^n &= -i\frac{e^2}{16\Lambda^2} \tilde{a}_n \epsilon_{ijk} W_{\mu\alpha}^{(i)} W_\nu^{(j)} W^{(k)\alpha} \tilde{F}^{\mu\nu} \end{aligned}$$

where  $F, W$  are photon and  $W$  fields,  $L_6^0$  and  $L_6^c$  conserve  $C, P$  separately ( $\tilde{L}_6^0$  conserves only  $C$ ) and generate anomalous  $W^+W^-\gamma\gamma$  and  $ZZ\gamma\gamma$  couplings,  $L_6^n$  violates  $CP$  ( $\tilde{L}_6^n$  violates both  $C$  and  $P$ ) and generates an anomalous  $W^+W^-Z\gamma$  coupling, and  $\Lambda$  is an energy scale for new physics. For the  $ZZ\gamma\gamma$  coupling the  $CP$ -violating term represented by  $L_6^n$  does not contribute. These couplings are assumed to be real and to vanish at tree level in the Standard Model.

Within the same framework as above, a more recent description of the quartic couplings [3] treats the anomalous parts of the  $WW\gamma\gamma$  and  $ZZ\gamma\gamma$  couplings separately, leading to two sets parametrized as  $a_0^V/\Lambda^2$  and  $a_c^V/\Lambda^2$ , where  $V = W$  or  $Z$ .

With the discovery of a Higgs at the LHC in 2012, it is then useful to go to the next higher dimensional representation (dimension 8 operators) in which the gauge symmetry is broken by the conventional Higgs scalar doublet [3,4]. There are 14 operators which can contribute to the anomalous quartic coupling signal. Some of the operators have analogues in the dimension 6 scheme. The CMS collaboration, [5], have used this parametrization, in which the connections between the two schemes are also summarized:

$$\begin{aligned} \mathcal{L}_{AQGC} &= -\frac{e^2}{8} \frac{a_0^W}{\Lambda^2} F_{\mu\nu} F^{\mu\nu} W^{+a} W_a^- \\ &\quad -\frac{e^2}{16} \frac{a_c^W}{\Lambda^2} F_{\mu\nu} F^{\mu\alpha} (W^{+\nu} W_a^- + W^{-\nu} W_a^+) \\ &\quad -e^2 g^2 \frac{\kappa_0^W}{\Lambda^2} F_{\mu\nu} Z^{\mu\nu} W^{+a} W_a^- \end{aligned}$$

$$\begin{aligned} &\quad -\frac{e^2 g^2}{2} \frac{\kappa_c^W}{\Lambda^2} F_{\mu\nu} Z^{\mu\alpha} (W^{+\nu} W_a^- + W^{-\nu} W_a^+) \\ &\quad + \frac{f_{T,0}}{\Lambda^4} Tr[\widehat{W}_{\mu\nu} \widehat{W}^{\mu\nu}] \times Tr[\widehat{W}_{\alpha\beta} \widehat{W}^{\alpha\beta}] \end{aligned}$$

The energy scale of possible new physics is  $\Lambda$ , and  $g = e/\sin(\theta_W)$ ,  $e$  being the unit electric charge and  $\theta_W$  the Weinberg angle. The field tensors are described in [3,4].

The two dimension 6 operators  $a_0^W/\Lambda^2$  and  $a_c^W/\Lambda^2$  are associated with the  $WW\gamma\gamma$  vertex. Among dimension 8 operators,  $\kappa_0^W/\Lambda^2$  and  $\kappa_c^W/\Lambda^2$  are associated with the  $WWZ\gamma$  vertex, whereas the parameter  $f_{T,0}/\Lambda^4$  contributes to both vertices. There is a relationship between these two dimension 6 parameters and the dimension 8 parameters  $f_{M,i}/\Lambda^4$  as follows [3]:

$$\begin{aligned} \frac{a_0^W}{\Lambda^2} &= -\frac{4M_W^2}{g^2} \frac{f_{M,0}}{\Lambda^4} - \frac{8M_W^2}{g'^2} \frac{f_{M,2}}{\Lambda^4} \\ \frac{a_c^W}{\Lambda^2} &= -\frac{4M_W^2}{g^2} \frac{f_{M,1}}{\Lambda^4} - \frac{8M_W^2}{g'^2} \frac{f_{M,3}}{\Lambda^4} \end{aligned}$$

where  $g' = e/\cos(\theta_W)$  and  $M_W$  is the invariant mass of the  $W$  boson. This relation provides a translation between limits on dimension 6 operators  $a_{0,c}^W$  and  $f_{M,j}/\Lambda^4$ . It is further required [4] that  $f_{M,0} = 2f_{M,2}$  and  $f_{M,1} = 2f_{M,3}$  which suppresses contributions to the  $WWZ\gamma$  vertex. The complete set of Lagrangian contributions as presented in [4] corresponds to 19 anomalous couplings in total –  $f_{S,i}$ ,  $i = 1, 2$ ,  $f_{M,i}$ ,  $i = 0, \dots, 8$  and  $f_{T,i}$ ,  $i = 0, \dots, 9$  – each scaled by  $1/\Lambda^4$ .

The ATLAS collaboration [6], on the other hand, follows a K-matrix driven approach of Ref. 7 in which the anomalous couplings can be expressed in terms of two parameters  $\alpha_4$  and  $\alpha_5$ , which account for all BSM effects.

It is the early stages in the determination of quartic couplings by the LHC experiments. It is hoped that the two collaborations, ATLAS and CMS, will agree to use at least one common set of parameters to express these limits to enable the reader to make a comparison and allow for a possible LHC combination.

### References

1. G. Belanger and F. Boudjema, Phys. Lett. **B288**, 201 (1992).
2. J.W. Stirling and A. Werthenbach, Eur. Phys. J. **C14**, 103 (2000);  
J.W. Stirling and A. Werthenbach, Phys. Lett. **B466**, 369 (1999);  
A. Denner *et al.*, Eur. Phys. J. **C20**, 201 (2001);  
G. Montagna *et al.*, Phys. Lett. **B515**, 197 (2001).
3. G. Belanger *et al.*, Eur. Phys. J. **C13**, 283 (2000).
4. O.J.P.   boli, M.C. Gonzalez-Garcia, and S.M. Lietti, Phys. Rev. **D69**, 095005 (2004);  
O.J.P.   boli, M.C. Gonzalez-Garcia, and J.K. Mizukoshi, Phys. Rev. **D77**, 073005 (2006).
5. S. Chatrchyan *et al.*, Phys. Rev. **D90**, 032008 (2014);  
S. Chatrchyan *et al.*, Phys. Rev. Lett. **114**, 051801 (2015).
6. G. Aad *et al.*, Phys. Rev. Lett. **113**, 141803 (2014).

## 7. A. Albateanu, W. Killian, and J. Reuter, JHEP **0811**, 010 (2008).

### $a_0/\Lambda^2$ , $a_c/\Lambda^2$

Combining published and unpublished preliminary LEP results the following 95% CL intervals for the QGCs associated with the  $ZZ\gamma\gamma$  vertex are derived (CERN-PH-EP/2005-051 or hep-ex/0511027):

$$-0.008 < a_0^Z/\Lambda^2 < +0.021$$

$$-0.029 < a_c^Z/\Lambda^2 < +0.039$$

Anomalous Z quartic couplings can also be measured by the experiments at the Tevatron and the LHC. As discussed in the review on “Anomalous  $W/Z$  quartic couplings (QGCs),” the measurements are typically done using different operator expansions which then do not allow the results to be compared and averaged. At least one common framework should be agreed upon for use in future publications by the experiments.

VALUE	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		

<sup>1</sup> ABBIENDI 04L OPAL

<sup>2</sup> HEISTER 04A ALEP

<sup>3</sup> ACHARD 02G L3

- ABBIENDI 04L select  $20 e^+e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  acoplanar events in the energy range 180–209 GeV and  $176 e^+e^- \rightarrow q\bar{q}\gamma\gamma$  events in the energy range 130–209 GeV. These samples are used to constrain possible anomalous  $W^+W^-\gamma\gamma$  and  $ZZ\gamma\gamma$  quartic couplings. Further combining with the  $W^+W^-\gamma$  sample of ABBIENDI 04b the following one-parameter 95% CL limits are obtained:  $-0.007 < a_0^Z/\Lambda^2 < 0.023 \text{ GeV}^{-2}$ ,  $-0.029 < a_c^Z/\Lambda^2 < 0.029 \text{ GeV}^{-2}$ ,  $-0.020 < a_0^W/\Lambda^2 < 0.020 \text{ GeV}^{-2}$ ,  $-0.052 < a_c^W/\Lambda^2 < 0.037 \text{ GeV}^{-2}$ .
- In the CM energy range 183 to 209 GeV HEISTER 04A select  $30 e^+e^- \rightarrow \nu\bar{\nu}\gamma\gamma$  events with two acoplanar, high energy and high transverse momentum photons. The photon-photon acoplanarity is required to be  $> 5^\circ$ ,  $E_\gamma/\sqrt{s} > 0.025$  (the more energetic photon having energy  $> 0.2\sqrt{s}$ ),  $p_{T\gamma}/E_{\text{beam}} > 0.05$  and  $|\cos\theta_\gamma| < 0.94$ . A likelihood fit to the photon energy and recoil missing mass yields the following one-parameter 95% CL limits:  $-0.012 < a_0^Z/\Lambda^2 < 0.019 \text{ GeV}^{-2}$ ,  $-0.041 < a_c^Z/\Lambda^2 < 0.044 \text{ GeV}^{-2}$ ,  $-0.060 < a_0^W/\Lambda^2 < 0.055 \text{ GeV}^{-2}$ ,  $-0.099 < a_c^W/\Lambda^2 < 0.093 \text{ GeV}^{-2}$ .
- ACHARD 02G study  $e^+e^- \rightarrow Z\gamma\gamma \rightarrow q\bar{q}\gamma\gamma$  events using data at center-of-mass energies from 200 to 209 GeV. The photons are required to be isolated, each with energy  $> 5 \text{ GeV}$  and  $|\cos\theta| < 0.97$ , and the di-jet invariant mass to be compatible with that of the Z boson (74–111 GeV). Cuts on Z velocity ( $\beta < 0.73$ ) and on the energy of the most energetic photon reduce the backgrounds due to non-resonant production of the  $q\bar{q}\gamma\gamma$  state and due to ISR respectively, yielding a total of 40 candidate events of which 8.6 are expected to be due to background. The energy spectra of the least energetic photon are fitted for all ten center-of-mass energy values from 130 GeV to 209 GeV (as obtained adding to the present analysis 130–202 GeV data of ACCIARRI 01E, for a total of 137 events with an expected background of 34.1 events) to obtain the fitted values  $a_0/\Lambda^2 = 0.00 \pm 0.02 \text{ GeV}^{-2}$  and  $a_c/\Lambda^2 = 0.03 \pm 0.02 \text{ GeV}^{-2}$ , where the other parameter is kept fixed to its Standard Model value (0). A simultaneous fit to both parameters yields the 95% CL limits  $-0.02 \text{ GeV}^{-2} < a_0/\Lambda^2 < 0.03 \text{ GeV}^{-2}$  and  $-0.07 \text{ GeV}^{-2} < a_c/\Lambda^2 < 0.05 \text{ GeV}^{-2}$ .

## Z REFERENCES

AAD 15BT JHEP 1509 049	G. Aad et al. (ATLAS Collab.)
AAD 15I PRL 114 121801	G. Aad et al. (ATLAS Collab.)
KHACHATRYAN... 15AC JHEP 1504 164	V. Khachatryan et al. (CMS Collab.)
KHACHATRYAN... 15B PL B740 250	V. Khachatryan et al. (CMS Collab.)
KHACHATRYAN... 15B EPJ C75 511	V. Khachatryan et al. (CMS Collab.)
AAD 14AU PR D90 072010	G. Aad et al. (ATLAS Collab.)
AAD 14N PRL 112 231806	G. Aad et al. (ATLAS Collab.)
AALTONEN 14E PRL 112 111803	T. Aaltonen et al. (CDF Collab.)
CHATRCHYAN 14AB PR D89 092005	S. Chatrchyan et al. (CMS Collab.)
AAD 13AN PR D87 112003	G. Aad et al. (ATLAS Collab.)
Also PR D91 119901 (err.)	G. Aad et al. (ATLAS Collab.)
AAD 13Z JHEP 1303 128	G. Aad et al. (ATLAS Collab.)
CHATRCHYAN 13B JHEP 1301 063	S. Chatrchyan et al. (CMS Collab.)
CHATRCHYAN 13BI JHEP 1310 164	S. Chatrchyan et al. (CMS Collab.)
SCHAEF 13A PRPL 532 119	S. Schaefer et al. (ALEPH Collab., DELPHI, L3+ (ATLAS Collab.))
AAD 12BX PL B717 45	G. Aad et al. (CMS Collab.)
ABAZOV 12S PR D85 052001	V.M. Abazov et al. (CMS Collab.)
CHATRCHYAN 12BM JHEP 1212 034	S. Chatrchyan et al. (CMS Collab.)
AALTONEN 11S PRL 107 051802	T. Aaltonen et al. (CDF Collab.)
ABAZOV 11D PR D84 012007	V.M. Abazov et al. (DO Collab.)
CHATRCHYAN 11M PL B701 535	S. Chatrchyan et al. (CMS Collab.)
ABAZOV 09L PRL 102 201802	V.M. Abazov et al. (DO Collab.)
BEDDALL 09 PL B670 300	A. Beddall, A. Beddall, A. Bingul (UGAZ Collab.)
SCHAEF 09 JHEP 0904 124	S. Schaefer et al. (ALEPH Collab.)
ABAZOV 08K PRL 100 131801	V.M. Abazov et al. (DO Collab.)
ABAZOV 07M PL B653 378	V.M. Abazov et al. (DO Collab.)
ABDALLAH 07C EPJ C51 525	J. Abdallah et al. (DELPHI Collab.)
ABDALLAH 06E PL B639 179	J. Abdallah et al. (DELPHI Collab.)
AKTAS 06 PL B632 35	A. Aktas et al. (HI Collab.)
LEP-SLC 06 PRPL 427 257	ALEPH, DELPHI, L3, OPAL, SLD and working groups
SCHAEF 06A PL B639 192	S. Schaefer et al. (ALEPH Collab.)
ABDALLAH 05 EPJ C40 1	J. Abdallah et al. (DELPHI Collab.)
ABDALLAH 05C EPJ C44 299	J. Abdallah et al. (DELPHI Collab.)
ABE 05 PRL 94 091801	K. Abe et al. (SLD Collab.)
ABE 05F PR D71 112004	K. Abe et al. (SLD Collab.)
ACOSTA 05M PR D71 052002	D. Acosta et al. (CDF Collab.)
ABBIENDI 04B PL B580 17	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 04C EPJ C32 303	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 04E PL B586 167	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 04G EPJ C33 173	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 04L PR D70 032005	G. Abbiendi et al. (OPAL Collab.)
ABDALLAH 04F EPJ C34 109	J. Abdallah et al. (DELPHI Collab.)
ABE 04C PR D69 072003	K. Abe et al. (SLD Collab.)
ACHARD 04C PL B585 42	P. Achard et al. (L3 Collab.)
ACHARD 04H PL B597 119	P. Achard et al. (L3 Collab.)
HEISTER 04A PL B602 31	A. Heister et al. (ALEPH Collab.)
ABBIENDI 03P PL B577 18	G. Abbiendi et al. (OPAL Collab.)
ABDALLAH 03H PL B569 129	J. Abdallah et al. (DELPHI Collab.)
ABDALLAH 03K PL B576 29	J. Abdallah et al. (DELPHI Collab.)
ABE 03F PRL 90 141804	K. Abe et al. (SLD Collab.)
ACHARD 03D PL B572 133	P. Achard et al. (L3 Collab.)
ACHARD 03G PL B577 109	P. Achard et al. (L3 Collab.)
ABBIENDI 02I PL B546 29	G. Abbiendi et al. (OPAL Collab.)
ABE 02G PRL 88 151801	K. Abe et al. (SLD Collab.)
ACHARD 02G PL B540 43	P. Achard et al. (L3 Collab.)
HEISTER 02B PL B526 34	A. Heister et al. (ALEPH Collab.)
HEISTER 02C PL B528 19	A. Heister et al. (ALEPH Collab.)
HEISTER 02H EPJ C24 177	A. Heister et al. (ALEPH Collab.)
ABBIENDI 01A EPJ C19 587	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 01G EPJ C18 447	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 01K PL B516 1	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 01N EPJ C20 445	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 01O EPJ C21 1	G. Abbiendi et al. (OPAL Collab.)
ABE 01B PRL 86 1162	K. Abe et al. (SLD Collab.)
ACCIARRI 01C PL B505 47	M. Acciarri et al. (L3 Collab.)
ACCIARRI 01I PL B497 23	M. Acciarri et al. (L3 Collab.)
HEISTER 01 EPJ C20 401	A. Heister et al. (ALEPH Collab.)
HEISTER 01D EPJ C22 201	A. Heister et al. (ALEPH Collab.)
ABBIENDI 00N PL B476 256	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI,G 00C EPJ C17 553	G. Abbiendi et al. (OPAL Collab.)
ABE 00B PRL 84 5945	K. Abe et al. (SLD Collab.)
ABE 00D PRL 85 5059	K. Abe et al. (SLD Collab.)
ABREU 00 EPJ C12 225	P. Abreu et al. (DELPHI Collab.)
ABREU 00B EPJ C14 613	P. Abreu et al. (DELPHI Collab.)
ABREU 00E EPJ C14 585	P. Abreu et al. (DELPHI Collab.)
ABREU 00F EPJ C16 371	P. Abreu et al. (DELPHI Collab.)
ABREU 00P PL B475 429	P. Abreu et al. (DELPHI Collab.)
ACCIARRI 00 EPJ C13 47	M. Acciarri et al. (L3 Collab.)
ACCIARRI 00C EPJ C16 1	M. Acciarri et al. (L3 Collab.)
ACCIARRI 00J PL B479 79	M. Acciarri et al. (L3 Collab.)
ACCIARRI 00Q PL B489 93	M. Acciarri et al. (L3 Collab.)
BARATE 00B EPJ C16 597	R. Barate et al. (ALEPH Collab.)
BARATE 00C EPJ C14 1	R. Barate et al. (ALEPH Collab.)
BARATE 00D EPJ C16 613	R. Barate et al. (ALEPH Collab.)
ABBIENDI 99B EPJ C8 217	G. Abbiendi et al. (OPAL Collab.)
ABBIENDI 99I PL B447 157	G. Abbiendi et al. (OPAL Collab.)
ABE 99E PR D59 052001	K. Abe et al. (SLD Collab.)
ABE 99L PRL 83 1902	K. Abe et al. (SLD Collab.)
ABREU 99 EPJ C6 19	P. Abreu et al. (DELPHI Collab.)
ABREU 99B EPJ C10 415	P. Abreu et al. (DELPHI Collab.)
ABREU 99J PL B449 364	P. Abreu et al. (DELPHI Collab.)
ABREU 99U PL B462 425	P. Abreu et al. (DELPHI Collab.)
ABREU 99Y EPJ C10 219	P. Abreu et al. (DELPHI Collab.)
ACCIARRI 99D PL B448 152	M. Acciarri et al. (L3 Collab.)
ACCIARRI 99F PL B453 94	M. Acciarri et al. (L3 Collab.)
ACCIARRI 99G PL B450 281	M. Acciarri et al. (L3 Collab.)
ACCIARRI 99O PL B455 363	M. Acciarri et al. (L3 Collab.)
ABBOTT 98M PR D57 R3817	B. Abbott et al. (DO Collab.)
ABE 98D PRL 80 660	K. Abe et al. (SLD Collab.)
ABE 98I PRL 81 942	K. Abe et al. (SLD Collab.)
ABREU 98K PL B423 194	P. Abreu et al. (DELPHI Collab.)
ABREU 98L EPJ C5 585	P. Abreu et al. (DELPHI Collab.)
ACCIARRI 98G PL B431 199	M. Acciarri et al. (L3 Collab.)
ACCIARRI 98H PL B429 387	M. Acciarri et al. (L3 Collab.)
ACCIARRI 98U PL B439 225	M. Acciarri et al. (L3 Collab.)
ACKERSTAFF 98A EPJ C5 411	K. Ackerstaff et al. (OPAL Collab.)
ACKERSTAFF 98B EPJ C1 439	K. Ackerstaff et al. (OPAL Collab.)
ACKERSTAFF 98C PL B420 157	K. Ackerstaff et al. (OPAL Collab.)
ACKERSTAFF 98Q EPJ C4 19	K. Ackerstaff et al. (OPAL Collab.)
BARATE 98O PL B434 415	R. Barate et al. (ALEPH Collab.)
BARATE 98T EPJ C4 557	R. Barate et al. (ALEPH Collab.)
BARATE 98V EPJ C5 205	R. Barate et al. (ALEPH Collab.)
ABE 97 PRL 78 17	K. Abe et al. (SLD Collab.)
ABREU 97C ZPHY C73 243	P. Abreu et al. (DELPHI Collab.)
ABREU 97E PL B398 207	P. Abreu et al. (DELPHI Collab.)
ABREU 97D PL B404 194	P. Abreu et al. (DELPHI Collab.)
ACCIARRI 97D PL B393 465	M. Acciarri et al. (L3 Collab.)
ACCIARRI 97J PL B407 351	M. Acciarri et al. (L3 Collab.)
ACCIARRI 97L PL B407 389	M. Acciarri et al. (L3 Collab.)
ACCIARRI 97P PL B413 167	M. Acciarri et al. (L3 Collab.)
ACKERSTAFF 97M ZPHY C74 413	K. Ackerstaff et al. (OPAL Collab.)
ACKERSTAFF 97S PL B412 210	K. Ackerstaff et al. (OPAL Collab.)
ACKERSTAFF 97T ZPHY C76 387	K. Ackerstaff et al. (OPAL Collab.)
ACKERSTAFF 97W ZPHY C76 425	K. Ackerstaff et al. (OPAL Collab.)
ALEXANDER 97C ZPHY C73 379	G. Alexander et al. (OPAL Collab.)
ALEXANDER 97D ZPHY C73 569	G. Alexander et al. (OPAL Collab.)
ALEXANDER 97E ZPHY C73 587	G. Alexander et al. (OPAL Collab.)
BARATE 97D PL B405 191	R. Barate et al. (ALEPH Collab.)
BARATE 97E PL B401 150	R. Barate et al. (ALEPH Collab.)
BARATE 97F PL B401 163	R. Barate et al. (ALEPH Collab.)
BARATE 97H PL B402 213	R. Barate et al. (ALEPH Collab.)
BARATE 97J ZPHY C74 451	R. Barate et al. (ALEPH Collab.)
ABREU 96R ZPHY C72 31	P. Abreu et al. (DELPHI Collab.)
ABREU 96S PL B389 405	P. Abreu et al. (DELPHI Collab.)
ABREU 96U ZPHY C73 61	P. Abreu et al. (DELPHI Collab.)
ACCIARRI 96 PL B371 126	M. Acciarri et al. (L3 Collab.)
ADAM 96 ZPHY C69 561	W. Adam et al. (DELPHI Collab.)
ADAM 96B ZPHY C70 371	W. Adam et al. (DELPHI Collab.)
ALEXANDER 96B ZPHY C70 197	G. Alexander et al. (OPAL Collab.)
ALEXANDER 96F PL B370 185	G. Alexander et al. (OPAL Collab.)
ALEXANDER 96N PL B384 343	G. Alexander et al. (OPAL Collab.)
ALEXANDER 96R ZPHY C72 1	G. Alexander et al. (OPAL Collab.)
BUSKULIC 96D ZPHY C69 393	D. Buskulic et al. (ALEPH Collab.)
BUSKULIC 96H ZPHY C69 379	D. Buskulic et al. (ALEPH Collab.)
BUSKULIC 96T PL B384 449	D. Buskulic et al. (ALEPH Collab.)
BUSKULIC 96Y PL B388 648	D. Buskulic et al. (ALEPH Collab.)
ABE 95J PRL 74 2880	K. Abe et al. (SLD Collab.)
ABREU 95 ZPHY C65 709 (erratum)	P. Abreu et al. (DELPHI Collab.)
ABREU 95D ZPHY C66 323	P. Abreu et al. (DELPHI Collab.)
ABREU 95L ZPHY C65 587	P. Abreu et al. (DELPHI Collab.)
ABREU 95M ZPHY C65 603	P. Abreu et al. (DELPHI Collab.)
ABREU 95O ZPHY C67 543	P. Abreu et al. (DELPHI Collab.)
ABREU 95R ZPHY C68 353	P. Abreu et al. (DELPHI Collab.)
ABREU 95V ZPHY C68 541	P. Abreu et al. (DELPHI Collab.)
ABREU 95W PL B361 207	P. Abreu et al. (DELPHI Collab.)
ABREU 95X ZPHY C69 1	P. Abreu et al. (DELPHI Collab.)
ACCIARRI 95B PL B345 589	M. Acciarri et al. (L3 Collab.)
ACCIARRI 95C PL B345 609	M. Acciarri et al. (L3 Collab.)
ACCIARRI 95D PL B353 136	M. Acciarri et al. (L3 Collab.)
AKERS 95C ZPHY C65 47	R. Akers et al. (OPAL Collab.)
AKERS 95R ZPHY C67 389	R. Akers et al. (OPAL Collab.)
AKERS 95W ZPHY C67 555	R. Akers et al. (OPAL Collab.)
AKERS 95X ZPHY C68 1	R. Akers et al. (OPAL Collab.)
AKERS 95Z ZPHY C68 203	R. Akers et al. (OPAL Collab.)
ALEXANDER 95D PL B358 162	G. Alexander et al. (OPAL Collab.)

# Gauge & Higgs Boson Particle Listings

## Z, H<sup>0</sup>

BUSKULIC	95R	ZPHY C69 15	D. Buskulis <i>et al.</i>	(ALEPH Collab.)
MIYABAYASHI	95	PL B347 171	K. Miyabayashi <i>et al.</i>	(TOPAZ Collab.)
ABE	94C	PRL 73 25	K. Abe <i>et al.</i>	(SLD Collab.)
ABREU	94B	PL B327 386	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	94P	PL B341 109	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	94P	ZPHY C63 181	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	94G	ZPHY C62 179	D. Buskulis <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94J	ZPHY C62 1	D. Buskulis <i>et al.</i>	(ALEPH Collab.)
VILAIN	94	PL B320 203	P. Vilain <i>et al.</i>	(CHARM II Collab.)
ABREU	93	PL B298 236	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	93I	ZPHY C59 533	P. Abreu <i>et al.</i>	(DELPHI Collab.)
Also		ZPHY C65 709 (erratum)	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	93L	PL B318 249	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	93	PL B305 407	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	93D	ZPHY C58 219	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	93E	PL B311 391	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ADRIANI	93	PL B301 136	O. Adriani <i>et al.</i>	(L3 Collab.)
ADRIANI	93I	PL B316 427	O. Adriani <i>et al.</i>	(L3 Collab.)
BUSKULIC	93L	PL B313 520	D. Buskulis <i>et al.</i>	(ALEPH Collab.)
NOVIKOV	93C	PL B298 453	V.A. Novikov, L.B. Okun, M.I. Vysotsky	(ITEP Collab.)
ABREU	92I	PL B277 371	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	92M	PL B289 199	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	92B	ZPHY C53 539	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	92L	PL B294 436	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	92N	PL B295 357	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ADEVA	92	PL B275 209	B. Adeva <i>et al.</i>	(L3 Collab.)
ADRIANI	92D	PL B292 454	O. Adriani <i>et al.</i>	(L3 Collab.)
ALITTI	92B	PL B276 354	J. Alitti <i>et al.</i>	(UA2 Collab.)
BUSKULIC	92D	PL B292 210	D. Buskulis <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	92E	PL B294 145	D. Buskulis <i>et al.</i>	(ALEPH Collab.)
DECAMP	92	PRPL 216 253	D. Decamp <i>et al.</i>	(ALEPH Collab.)
ABE	91E	PRL 67 1502	F. Abe <i>et al.</i>	(CDF Collab.)
ABREU	91H	ZPHY C50 185	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	91B	PL B273 338	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ADACHI	91	PL B255 613	I. Adachi <i>et al.</i>	(TOPAZ Collab.)
ADEVA	91I	PL B259 199	B. Adeva <i>et al.</i>	(L3 Collab.)
AKRAWY	91F	PL B257 531	M.Z. Akrawy <i>et al.</i>	(OPAL Collab.)
DECAMP	91B	PL B259 377	D. Decamp <i>et al.</i>	(ALEPH Collab.)
DECAMP	91J	PL B266 218	D. Decamp <i>et al.</i>	(ALEPH Collab.)
JACOBSEN	91	PRL 67 3347	R.G. Jacobsen <i>et al.</i>	(Mark II Collab.)
SHIMONAKA	91	PL B268 457	A. Shimonaka <i>et al.</i>	(TOPAZ Collab.)
ABE	90I	ZPHY C48 13	K. Abe <i>et al.</i>	(VENUS Collab.)
ABRAMS	90	PRL 64 1334	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
AKRAWY	90J	PL B246 285	M.Z. Akrawy <i>et al.</i>	(OPAL Collab.)
BEHREND	90D	ZPHY C47 333	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BRAUNSCH...	90	ZPHY C48 433	W. Braunschweig <i>et al.</i>	(TASSO Collab.)
ELSEN	90	ZPHY C46 349	E. Elsen <i>et al.</i>	(JADE Collab.)
HEGNER	90	ZPHY C46 547	S. Hegner <i>et al.</i>	(JADE Collab.)
STUART	90	PRL 64 983	D. Stuart <i>et al.</i>	(AMY Collab.)
ABE	89	PRL 62 613	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	89C	PRL 63 720	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	89L	PL B232 425	K. Abe <i>et al.</i>	(VENUS Collab.)
ABRAMS	89B	PRL 63 2173	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
ABRAMS	89D	PL 63 2780	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
ALBAJAR	89	ZPHY C44 15	C. Albajar <i>et al.</i>	(UA1 Collab.)
BACALA	89	PL B218 112	A. Bacala <i>et al.</i>	(AMY Collab.)
BAND	89	PL B218 369	H.R. Band <i>et al.</i>	(MAC Collab.)
GREENSHAW	89	ZPHY C42 1	T. Greenshaw <i>et al.</i>	(JADE Collab.)
OULD-SAADA	89	ZPHY C44 567	F. Ould-Saada <i>et al.</i>	(JADE Collab.)
SAGAWA	89	PRL 63 2341	H. Sagawa <i>et al.</i>	(AMY Collab.)
ADACHI	88C	PL B208 319	I. Adachi <i>et al.</i>	(TOPAZ Collab.)
ADEVA	88	PR D38 2665	B. Adeva <i>et al.</i>	(Mark-J Collab.)
BRAUNSCH...	88D	ZPHY C40 163	W. Braunschweig <i>et al.</i>	(TASSO Collab.)
ANSARI	87	PL B186 440	R. Ansari <i>et al.</i>	(UA2 Collab.)
BEHREND	87C	PL B191 209	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BARTEL	86C	ZPHY C30 371	W. Bartel <i>et al.</i>	(JADE Collab.)
Also		ZPHY C26 507	W. Bartel <i>et al.</i>	(JADE Collab.)
Also		PL 108B 140	W. Bartel <i>et al.</i>	(JADE Collab.)
ASH	85	PRL 55 1831	W.W. Ash <i>et al.</i>	(MAC Collab.)
BARTEL	85F	PL 161B 188	W. Bartel <i>et al.</i>	(JADE Collab.)
DERRICK	85	PR D31 2352	M. Derrick <i>et al.</i>	(HRS Collab.)
FERNANDEZ	85	PRL 54 1624	E. Fernandez <i>et al.</i>	(MAC Collab.)
LEVI	83	PRL 51 1941	M.E. Levi <i>et al.</i>	(Mark II Collab.)
BEHREND	82	PL 114B 282	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BRANDELIK	82C	PL 110B 173	R. Brandelik <i>et al.</i>	(TASSO Collab.)

## H<sup>0</sup>

$J = 0$

In the following  $H^0$  refers to the signal that has been discovered in the Higgs searches. Whereas the observed signal is labeled as a spin 0 particle and is called a Higgs Boson, the detailed properties of  $H^0$  and its role in the context of electroweak symmetry breaking need to be further clarified. These issues are addressed by the measurements listed below.

Concerning mass limits and cross section limits that have been obtained in the searches for neutral and charged Higgs bosons, see the sections “Searches for Neutral Higgs Bosons” and “Searches for Charged Higgs Bosons ( $H^\pm$  and  $H^{\pm\pm}$ )”, respectively.

### H<sup>0</sup> MASS

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b>125.09±0.21±0.11</b>	1,2 AAD	15B LHC	$pp$ , 7, 8 TeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
125.07±0.25±0.14	2 AAD	15B LHC	$pp$ , 7, 8 TeV, $\gamma\gamma$
125.15±0.37±0.15	2 AAD	15B LHC	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
126.02±0.43±0.27	AAD	15B ATLS	$pp$ , 7, 8 TeV, $\gamma\gamma$
124.51±0.52±0.04	AAD	15B ATLS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
125.59±0.42±0.17	AAD	15B CMS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
125.02 $^{+0.26}_{-0.27}$ ±0.15	3 KHACHATRY...15AM	CMS	$pp$ , 7, 8 TeV
125.36±0.37±0.18	1,4 AAD	14W ATLS	$pp$ , 7, 8 TeV
125.98±0.42±0.28	4 AAD	14W ATLS	$pp$ , 7, 8 TeV, $\gamma\gamma$
124.51±0.52±0.06	4 AAD	14W ATLS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
125.6 ±0.4 ±0.2	5 CHATRCHYAN14AA	CMS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$

122 ±7	6 CHATRCHYAN14K	CMS	$pp$ , 7, 8 TeV, $\tau\tau$
124.70±0.31±0.15	7 KHACHATRY...14P	CMS	$pp$ , 7, 8 TeV, $\gamma\gamma$
125.5 ±0.2 $^{+0.5}_{-0.6}$	1,8 AAD	13AK ATLS	$pp$ , 7, 8 TeV
126.8 ±0.2 ±0.7	8 AAD	13AK ATLS	$pp$ , 7, 8 TeV, $\gamma\gamma$
124.3 $^{+0.6}_{-0.5}$ ±0.5	8 AAD	13AK ATLS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
125.8 ±0.4 ±0.4	1,9 CHATRCHYAN13J	CMS	$pp$ , 7, 8 TeV
126.2 ±0.6 ±0.2	9 CHATRCHYAN13J	CMS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
126.0 ±0.4 ±0.4	1,10 AAD	12AI ATLS	$pp$ , 7, 8 TeV
125.3 ±0.4 ±0.5	1,11 CHATRCHYAN12N	CMS	$pp$ , 7, 8 TeV

- 1 Combined value from  $\gamma\gamma$  and  $ZZ^* \rightarrow 4\ell$  final states.
- 2 ATLAS and CMS data are fitted simultaneously.
- 3 KHACHATRYAN 15AM use up to 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV.
- 4 AAD 14W use 4.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at 8 TeV.
- 5 CHATRCHYAN 14AA use 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV.
- 6 CHATRCHYAN 14K use 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV.
- 7 KHACHATRYAN 14P use 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV.
- 8 AAD 13AK use 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. Superseded by AAD 14W.
- 9 CHATRCHYAN 13J use 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 12.2 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV.
- 10 AAD 12AI obtain results based on 4.6–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.8–5.9 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. An excess of events over background with a local significance of 5.9  $\sigma$  is observed at  $m_{H^0} = 126$  GeV. See also AAD 12DA.
- 11 CHATRCHYAN 12N obtain results based on 4.9–5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.1–5.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. An excess of events over background with a local significance of 5.0  $\sigma$  is observed at about  $m_{H^0} = 125$  GeV. See also CHATRCHYAN 12BY and CHATRCHYAN 13Y.

### H<sup>0</sup> SPIN AND CP PROPERTIES

The observation of the signal in the  $\gamma\gamma$  final state rules out the possibility that the discovered particle has spin 1, as a consequence of the Landau-Yang theorem. This argument relies on the assumptions that the decaying particle is an on-shell resonance and that the decay products are indeed two photons rather than two pairs of boosted photons, which each could in principle be misidentified as a single photon.

Concerning distinguishing the spin 0 hypothesis from a spin 2 hypothesis, some care has to be taken in modelling the latter in order to ensure that the discriminating power is actually based on the spin properties rather than on unphysical behavior that may affect the model of the spin 2 state.

Under the assumption that the observed signal consists of a single state rather than an overlap of more than one resonance, it is sufficient to discriminate between distinct hypotheses in the spin analyses. On the other hand, the determination of the  $CP$  properties is in general much more difficult since in principle the observed state could consist of any admixture of  $CP$ -even and  $CP$ -odd components. As a first step, the compatibility of the data with distinct hypotheses of pure  $CP$ -even and pure  $CP$ -odd states with different spin assignments has been investigated. In order to treat the case of a possible mixing of different  $CP$  states, certain cross section ratios are considered. Those cross section ratios need to be distinguished from the amount of mixing between a  $CP$ -even and a  $CP$ -odd state, as the cross section ratios depend in addition also on the coupling strengths of the  $CP$ -even and  $CP$ -odd components to the involved particles. A small relative coupling implies a small sensitivity of the corresponding cross section ratio to effects of  $CP$  mixing.

VALUE	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
1 AAD	16 ATLS	$H^0 \rightarrow \gamma\gamma$	
2 AAD	15AX ATLS	$H^0 \rightarrow WW^*$	
3 AAD	15CI ATLS	$H^0 \rightarrow ZZ^*, WW^*, \gamma\gamma$	
4 AALTONEN	15 TEVA	$p\bar{p} \rightarrow WH^0, ZH^0, H^0 \rightarrow b\bar{b}$	
5 AALTONEN	15B CDF	$p\bar{p} \rightarrow WH^0, ZH^0, H^0 \rightarrow b\bar{b}$	
6 KHACHATRY...15Y	CMS	$H^0 \rightarrow 4\ell, WW^*, \gamma\gamma$	
7 ABZOV	14F D0	$p\bar{p} \rightarrow WH^0, ZH^0, H^0 \rightarrow b\bar{b}$	
8 CHATRCHYAN14AA	CMS	$H^0 \rightarrow ZZ^*$	
9 CHATRCHYAN14G	CMS	$H^0 \rightarrow WW^*$	
10 KHACHATRY...14P	CMS	$H^0 \rightarrow \gamma\gamma$	
11 AAD	13AJ ATLS	$H^0 \rightarrow \gamma\gamma, ZZ^* \rightarrow 4\ell, WW^* \rightarrow \ell\nu\ell\nu$	
12 CHATRCHYAN13J	CMS	$H^0 \rightarrow ZZ^* \rightarrow 4\ell$	

- 1 AAD 16 study  $H^0 \rightarrow \gamma\gamma$  with an effective Lagrangian including  $CP$  even and odd terms in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. The data is consistent with the expectations for the Higgs boson of the Standard Model. Limits on anomalous couplings are also given.
- 2 AAD 15AX compare the  $J^{CP} = 0^+$  Standard Model assignment with other  $J^{CP}$  hypotheses in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV, using the process  $H^0 \rightarrow WW^* \rightarrow e\nu\mu\nu$ .  $2^+$  hypotheses are excluded at 84.5–99.4%CL,  $0^-$  at 96.5%CL,  $0^+$  (field strength coupling) at 70.8%CL. See their Fig. 19 for limits on possible  $CP$  mixture parameters.



- <sup>3</sup> AAD 15c compare the  $J^{CP} = 0^+$  Standard Model assignment with other  $J^{CP}$  hypotheses in  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ , using the processes  $H^0 \rightarrow ZZ^* \rightarrow 4\ell$ ,  $H^0 \rightarrow \gamma\gamma$  and combine with AAD 15Ax data.  $0^+$  (field strength coupling),  $0^-$  and several  $2^+$  hypotheses are excluded at more than 99.9% CL. See their Tables 7–9 for limits on possible  $CP$  mixture parameters.
- <sup>4</sup> AALTONEN 15 combine AALTONEN 15b and ABAZOV 14F data. An upper limit of 0.36 of the Standard Model production rate at 95% CL is obtained both for a  $0^-$  and a  $2^+$  state. Assuming the SM event rate, the  $J^{CP} = 0^-$  ( $2^+$ ) hypothesis is excluded at the  $5.0\sigma$  ( $4.9\sigma$ ) level.
- <sup>5</sup> AALTONEN 15B compare the  $J^{CP} = 0^+$  Standard Model assignment with other  $J^{CP}$  hypotheses in  $9.45 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ , using the processes  $ZH^0 \rightarrow \ell\ell b\bar{b}$ ,  $WH^0 \rightarrow \ell\nu b\bar{b}$ , and  $ZH^0 \rightarrow \nu\nu b\bar{b}$ . Bounds on the production rates of  $0^-$  and  $2^+$  (graviton-like) states are set, see their tables II and III.
- <sup>6</sup> KHACHATRYAN 15V compare the  $J^{CP} = 0^+$  Standard Model assignment with other  $J^{CP}$  hypotheses in up to  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and up to  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ , using the processes  $H^0 \rightarrow 4\ell$ ,  $H^0 \rightarrow WW^*$ , and  $H^0 \rightarrow \gamma\gamma$ .  $0^-$  is excluded at 99.98% CL, and several  $2^+$  hypotheses are excluded at more than 99% CL. Spin 1 models are excluded at more than 99.99% CL in  $ZZ^*$  and  $WW^*$  modes. Limits on anomalous couplings and several cross section fractions, treating the case of  $CP$ -mixed states, are also given.
- <sup>7</sup> ABAZOV 14F compare the  $J^{CP} = 0^+$  Standard Model assignment with  $J^{CP} = 0^-$  and  $2^+$  (graviton-like coupling) hypotheses in up to  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . They use kinematic correlations between the decay products of the vector boson and the Higgs boson in the final states  $ZH \rightarrow \ell\ell b\bar{b}$ ,  $WH \rightarrow \ell\nu b\bar{b}$ , and  $ZH \rightarrow \nu\nu b\bar{b}$ . The  $0^-$  ( $2^+$ ) hypothesis is excluded at 97.6% CL (99.0% CL). In order to treat the case of a possible mixture of a  $0^+$  state with another  $J^{CP}$  state, the cross section fractions  $f_X = \sigma_X/(\sigma_0 + \sigma_X)$  are considered, where  $X = 0^-, 2^+$ . Values for  $f_0 - (f_2 +)$  above 0.80 (0.67) are excluded at 95% CL under the assumption that the total cross section is that of the SM Higgs boson.
- <sup>8</sup> CHATRCHYAN 14AA compare the  $J^{CP} = 0^+$  Standard Model assignment with various  $J^{CP}$  hypotheses in  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ .  $J^{CP} = 0^-$  and  $1^\pm$  hypotheses are excluded at 99% CL, and several  $J = 2$  hypotheses are excluded at 95% CL. In order to treat the case of a possible mixture of a  $0^+$  state with another  $J^{CP}$  state, the cross section fraction  $f_{a3} = |a_3|^2 \sigma_3 / (|a_1|^2 \sigma_1 + |a_2|^2 \sigma_2 + |a_3|^2 \sigma_3)$  is considered, where the case  $a_3 = 1$ ,  $a_1 = a_2 = 0$  corresponds to a pure  $CP$ -odd state. Assuming  $a_2 = 0$ , a value for  $f_{a3}$  above 0.51 is excluded at 95% CL.
- <sup>9</sup> CHATRCHYAN 14G compare the  $J^{CP} = 0^+$  Standard Model assignment with  $J^{CP} = 0^-$  and  $2^+$  (graviton-like coupling) hypotheses in  $4.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.4 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . Varying the fraction of the production of the  $2^+$  state via  $gg$  and  $q\bar{q}$ ,  $2^+$  hypotheses are disfavored at CL between 83.7 and 99.8%. The  $0^-$  hypothesis is disfavored against  $0^+$  at the 65.3% CL.
- <sup>10</sup> KHACHATRYAN 14P compare the  $J^{CP} = 0^+$  Standard Model assignment with a  $2^+$  (graviton-like coupling) hypothesis in  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . Varying the fraction of the production of the  $2^+$  state via  $gg$  and  $q\bar{q}$ ,  $2^+$  hypotheses are disfavored at CL between 71 and 94%.
- <sup>11</sup> AAD 13AJ compare the spin 0,  $CP$ -even hypothesis with specific alternative hypotheses of spin 0,  $CP$ -odd, spin 1,  $CP$ -even and  $CP$ -odd, and spin 2,  $CP$ -even modes using the Higgs boson decays  $H \rightarrow \gamma\gamma$ ,  $H \rightarrow ZZ^* \rightarrow 4\ell$  and  $H \rightarrow WW^* \rightarrow \ell\nu\ell\nu$  and combinations thereof. The data are compatible with the spin 0,  $CP$ -even hypothesis, while all other tested hypotheses are excluded at confidence levels above 97.8%.
- <sup>12</sup> CHATRCHYAN 13J study angular distributions of the lepton pairs in the  $ZZ^*$  channel where both  $Z$  bosons decay to  $e$  or  $\mu$  pairs. Under the assumption that the observed particle has spin 0, the data are found to be consistent with the pure  $CP$ -even hypothesis, while the pure  $CP$ -odd hypothesis is disfavored.

## $H^0$ DECAY WIDTH

The total decay width for a light Higgs boson with a mass in the observed range is not expected to be directly observable at the LHC. For the case of the Standard Model the prediction for the total width is about 4 MeV, which is three orders of magnitude smaller than the experimental mass resolution. There is no indication from the results observed so far that the natural width is broadened by new physics effects to such an extent that it could be directly observable. Furthermore, as all LHC Higgs channels rely on the identification of Higgs decay products, the total Higgs width cannot be measured indirectly without additional assumptions. The different dependence of on-peak and off-peak contributions on the total width in Higgs decays to  $ZZ^*$  and interference effects between signal and background in Higgs decays to  $\gamma\gamma$  can provide additional information in this context. Constraints on the total width from the combination of on-peak and off-peak contributions in Higgs decays to  $ZZ^*$  rely on the assumption of equal on- and off-shell effective couplings. Without an experimental determination of the total width or further theoretical assumptions, only ratios of couplings can be determined at the LHC rather than absolute values of couplings.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.7</b>	95	<sup>1</sup> KHACHATRYAN...15AM	CMS	$pp$ , 7, 8 TeV
>3.5	$\times 10^{-12}$	<sup>2</sup> KHACHATRYAN...15BA	CMS	$pp$ , 7, 8 TeV, flight distance
<5.0	95	<sup>3</sup> AAD	14W ATLS	$pp$ , 7, 8 TeV, $\gamma\gamma$
<2.6	95	<sup>3</sup> AAD	14W ATLS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.0227	95	<sup>4</sup> AAD	15BE ATLS	$pp$ , 8 TeV, $ZZ^*$ , $WW^*$
<0.046	95	<sup>5</sup> KHACHATRYAN...15BA	CMS	$pp$ , 7, 8 TeV, $ZZ^*$ $\rightarrow 4\ell$
<3.4	95	<sup>6</sup> CHATRCHYAN 14AA	CMS	$pp$ , 7, 8 TeV, $ZZ^* \rightarrow 4\ell$
<0.022	95	<sup>7</sup> KHACHATRYAN...14D	CMS	$pp$ , 7, 8 TeV, $ZZ^*$
<2.4	95	<sup>8</sup> KHACHATRYAN...14P	CMS	$pp$ , 7, 8 TeV, $\gamma\gamma$

- <sup>1</sup> KHACHATRYAN 15AM combine  $\gamma\gamma$  and  $ZZ^* \rightarrow 4\ell$  results. The expected limit is 2.3 GeV.
- <sup>2</sup> KHACHATRYAN 15BA derive a lower limit on the total width from an upper limit on the decay flight distance  $\tau < 1.9 \times 10^{-13} \text{ s}$ .  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at 8 TeV are used.
- <sup>3</sup> AAD 14W use  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at 8 TeV. The expected limit is 6.2 GeV.
- <sup>4</sup> AAD 15BE derive constraints on the total width from comparing  $ZZ^*$  and  $WW^*$  production via on-shell and off-shell  $H^0$  using  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8 \text{ TeV}$ . The K factor for the background processes is assumed to be equal to that for the signal.
- <sup>5</sup> KHACHATRYAN 15BA derive constraints on the total width from comparing  $ZZ^*$  production via on-shell and off-shell  $H^0$  with an unconstrained anomalous coupling.  $4\ell$  final states in  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$  are used.
- <sup>6</sup> CHATRCHYAN 14AA use  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The expected limit is 2.8 GeV.
- <sup>7</sup> KHACHATRYAN 14D derive constraints on the total width from comparing  $ZZ^*$  production via on-shell and off-shell  $H^0$ .  $4\ell$  and  $\ell\ell\nu\nu$  final states in  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$  are used.
- <sup>8</sup> KHACHATRYAN 14P use  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The expected limit is 3.1 GeV.

## $H^0$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $WW^*$		
$\Gamma_2$ $ZZ^*$		
$\Gamma_3$ $\gamma\gamma$		
$\Gamma_4$ $b\bar{b}$		
$\Gamma_5$ $e^+e^-$	$< 1.9 \times 10^{-3}$	95%
$\Gamma_6$ $\mu^+\mu^-$		
$\Gamma_7$ $\tau^+\tau^-$		
$\Gamma_8$ $Z\gamma$		
$\Gamma_9$ $J/\psi\gamma$	$< 1.5 \times 10^{-3}$	95%
$\Gamma_{10}$ $\mathcal{T}(1S)\gamma$	$< 1.3 \times 10^{-3}$	95%
$\Gamma_{11}$ $\mathcal{T}(2S)\gamma$	$< 1.9 \times 10^{-3}$	95%
$\Gamma_{12}$ $\mathcal{T}(3S)\gamma$	$< 1.3 \times 10^{-3}$	95%
$\Gamma_{13}$ $\mu\tau$	$< 1.51 \%$	95%
$\Gamma_{14}$ invisible	$< 58 \%$	95%

## $H^0$ BRANCHING RATIOS

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	$\Gamma_5/\Gamma$
<b>&lt;1.9 <math>\times 10^{-3}</math></b>	95	<sup>1</sup> KHACHATRYAN...15H	CMS	
<sup>1</sup> KHACHATRYAN 15H use $5.0 \text{ fb}^{-1}$ of $pp$ collisions at $E_{\text{cm}} = 7 \text{ TeV}$ and $19.7 \text{ fb}^{-1}$ at 8 TeV.				

$\Gamma(J/\psi\gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
<b>&lt;1.5 <math>\times 10^{-3}</math></b>	95	<sup>1</sup> KHACHATRYAN...16B	CMS	8 TeV	
<b>&lt;1.5 <math>\times 10^{-3}</math></b>	95	<sup>2</sup> AAD	15I	ATLS 8 TeV	
<sup>1</sup> KHACHATRYAN 16B use $19.7 \text{ fb}^{-1}$ of $pp$ collision data at 8 TeV.					
<sup>2</sup> AAD 15I use $19.7 \text{ fb}^{-1}$ of $pp$ collision data at 8 TeV.					

$\Gamma(\mathcal{T}(1S)\gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
<b>&lt;1.3 <math>\times 10^{-3}</math></b>	95	<sup>1</sup> AAD	15I	ATLS 8 TeV	
<sup>1</sup> AAD 15I use $19.7 \text{ fb}^{-1}$ of $pp$ collision data at 8 TeV.					

$\Gamma(\mathcal{T}(2S)\gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma$
<b>&lt;1.9 <math>\times 10^{-3}</math></b>	95	<sup>1</sup> AAD	15I	ATLS 8 TeV	
<sup>1</sup> AAD 15I use $19.7 \text{ fb}^{-1}$ of $pp$ collision data at 8 TeV.					

$\Gamma(\mathcal{T}(3S)\gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12}/\Gamma$
<b>&lt;1.3 <math>\times 10^{-3}</math></b>	95	<sup>1</sup> AAD	15I	ATLS 8 TeV	
<sup>1</sup> AAD 15I use $19.7 \text{ fb}^{-1}$ of $pp$ collision data at 8 TeV.					

$\Gamma(\mu\tau)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	$\Gamma_{13}/\Gamma$
<b>&lt;1.51 <math>\times 10^{-2}</math></b>	95	<sup>1</sup> KHACHATRYAN...15Q	CMS	
<sup>1</sup> KHACHATRYAN 15Q search for $H^0 \rightarrow \mu\tau$ with $\tau$ decaying electronically or hadronically in $19.7 \text{ fb}^{-1}$ of $pp$ collisions at $E_{\text{cm}} = 8 \text{ TeV}$ . The fit gives $B(H^0 \rightarrow \mu\tau) = (0.84^{+0.39}_{-0.37})\%$ with a significance of 2.4 $\sigma$ .				

# Gauge & Higgs Boson Particle Listings

## $H^0$

 $\Gamma(\text{invisible})/\Gamma_{\text{total}}$ 
 $\Gamma_{14}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.75	95	1 AAD	140 ATLS	$pp \rightarrow H^0 Z X$ , 7, 8 TeV
<0.58	95	2 CHATRCHYAN14B	CMS	$pp \rightarrow H^0 Z X$ , $qqH^0 X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.78	95	3 AAD	15BD ATLS	$pp \rightarrow H^0 W/Z X$ , 8 TeV
<0.81	95	4 CHATRCHYAN14B	CMS	$pp \rightarrow H^0 Z X$ , 7, 8 TeV
<0.65	95	5 CHATRCHYAN14B	CMS	$pp \rightarrow qqH^0 X$ , 8 TeV

- 1 AAD 140 search for  $pp \rightarrow H^0 Z X$ ,  $Z \rightarrow \ell\ell$ , with  $H^0$  decaying to invisible final states in 4.5 fb<sup>-1</sup> at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted limit on the branching ratio is given for  $m_{H^0} = 125.5$  GeV and assumes the Standard Model rate for  $H^0 Z$  production.
- 2 CHATRCHYAN 14B search for  $pp \rightarrow H^0 Z X$ ,  $Z \rightarrow \ell\ell$  and  $Z \rightarrow b\bar{b}$ , and also  $pp \rightarrow qqH^0 X$  with  $H^0$  decaying to invisible final states using data at  $E_{\text{cm}} = 7$  and 8 TeV. The quoted limit on the branching ratio is obtained from a combination of the limits from  $H^0 Z$  and  $qqH^0$ . It is given for  $m_{H^0} = 125$  GeV and assumes the Standard Model rates for the two production processes.
- 3 AAD 15BD search for  $pp \rightarrow H^0 WX$  and  $pp \rightarrow H^0 ZX$  with  $W$  or  $Z$  decaying hadronically and  $H^0$  decaying to invisible final states using data at  $E_{\text{cm}} = 8$  TeV. The quoted limit is given for  $m_{H^0} = 125$  GeV, assumes the Standard Model rates for the production processes and is based on a combination of the contributions from  $H^0 W$ ,  $H^0 Z$  and the gluon-fusion process.
- 4 CHATRCHYAN 14B search for  $pp \rightarrow H^0 ZX$  with  $H^0$  decaying to invisible final states and  $Z \rightarrow \ell\ell$  in 4.9 fb<sup>-1</sup> at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV, and also with  $Z \rightarrow b\bar{b}$  in 18.9 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted limit on the branching ratio is given for  $m_{H^0} = 125$  GeV and assumes the Standard Model rate for  $H^0 Z$  production.
- 5 CHATRCHYAN 14B search for  $pp \rightarrow qqH^0 X$  (vector boson fusion) with  $H^0$  decaying to invisible final states in 19.5 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted limit on the branching ratio is given for  $m_{H^0} = 125$  GeV and assumes the Standard Model rate for  $qqH^0$  production.

## $H^0$ SIGNAL STRENGTHS IN DIFFERENT CHANNELS

The  $H^0$  signal strength in a particular final state  $xx$  is given by the cross section times branching ratio in this channel normalized to the Standard Model (SM) value,  $\sigma \cdot \text{B}(H^0 \rightarrow xx) / (\sigma \cdot \text{B}(H^0 \rightarrow xx))_{\text{SM}}$ , for the specified mass value of  $H^0$ . For the SM predictions, see DITTMAYER 11, DITTMAYER 12, and HEINEMEYER 13A. Results for fiducial and differential cross sections are also listed below.

### Combined Final States

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.10±0.11 OUR AVERAGE</b>			
1.09±0.07±0.04±0.03±0.07 -0.06	1,2 AAD	16J LHC	$pp$ , 7, 8 TeV
1.44±0.59 -0.56	3 AALTONEN	13M TEVA	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.20±0.10±0.06±0.04±0.08 -0.07	2 AAD	16J ATLS	$pp$ , 7, 8 TeV
0.97±0.09±0.05±0.04±0.07 -0.03-0.06	2 AAD	16J CMS	$pp$ , 7, 8 TeV
1.18±0.10±0.07±0.08 -0.07	4 AAD	16K ATLS	$pp$ , 7, 8 TeV
0.75±0.28±0.13±0.08 -0.26-0.11-0.05	4 AAD	16K ATLS	$pp$ , 7 TeV
1.28±0.11±0.08±0.10 -0.07-0.08	4 AAD	16K ATLS	$pp$ , 8 TeV
	5 AAD	15P ATLS	$pp$ , 8 TeV, cross section
1.00±0.09±0.07±0.08 -0.07	6 KHACHATRYAN15AM	CMS	$pp$ , 7, 8 TeV
1.33±0.14 -0.10±0.15	7 AAD	13AK ATLS	$pp$ , 7 and 8 TeV
1.54±0.77 -0.73	8 AALTONEN	13L CDF	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
1.40±0.92 -0.88	9 ABAZOV	13L D0	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
1.4 ± 0.3	10 AAD	12AI ATLS	$pp \rightarrow H^0 X$ , 7, 8 TeV
1.2 ± 0.4	10 AAD	12AI ATLS	$pp \rightarrow H^0 X$ , 7 TeV
1.5 ± 0.4	10 AAD	12AI ATLS	$pp \rightarrow H^0 X$ , 8 TeV
0.87±0.23	11 CHATRCHYAN12N	CMS	$pp \rightarrow H^0 X$ , 7, 8 TeV

- 1 AAD 16J perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and 8 TeV. The signal strengths for individual production processes are  $1.03^{+0.16}_{-0.14}$  for gluon fusion,  $1.18^{+0.25}_{-0.23}$  for vector boson fusion,  $0.89^{+0.40}_{-0.38}$  for  $WH^0$  production,  $0.79^{+0.38}_{-0.36}$  for  $ZH^0$  production, and  $2.3^{+0.7}_{-0.6}$  for  $t\bar{t}H^0$  production.
- 2 The uncertainties represent statistics, experimental systematics, theory systematics on the background, and theory systematics on the signal. The quoted signal strengths are given for  $m_{H^0} = 125.09$  GeV. In the fit, relative branching ratios and relative production cross sections are fixed to those in the Standard Model.
- 3 AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations with up to 10.0 fb<sup>-1</sup> and 9.7 fb<sup>-1</sup>, respectively, of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.

- 4 AAD 16K use up to 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The third uncertainty in the measurement is theory systematics. The signal strengths for individual production modes are  $1.23^{+0.14}_{-0.08} \pm 0.09^{+0.16}_{-0.12}$  for gluon fusion,  $1.23^{+0.28}_{-0.27} \pm 0.13^{+0.11}_{-0.09}$  for vector boson fusion,  $0.80^{+0.31}_{-0.30} \pm 0.17^{+0.10}_{-0.05}$  for  $W/ZH^0$  production, and  $1.81^{+0.52}_{-0.50} \pm 0.58^{+0.31}_{-0.55} \pm 0.12$  for  $t\bar{t}H^0$  production. The quoted signal strengths are given for  $m_{H^0} = 125.36$  GeV.
- 5 AAD 15P measure total and differential cross sections of the process  $pp \rightarrow H^0 X$  at  $E_{\text{cm}} = 8$  TeV with 20.3 fb<sup>-1</sup>.  $\gamma\gamma$  and  $4\ell$  final states are used.  $\sigma(pp \rightarrow H^0 X) = 33.0 \pm 5.3 \pm 1.6$  pb is given. See their Figs. 2 and 3 for data on differential cross sections.
- 6 KHACHATRYAN 15AM use up to 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The third uncertainty in the measurement is theory systematics. Fits to each production mode give the value of  $0.85^{+0.19}_{-0.16}$  for gluon fusion,  $1.16^{+0.37}_{-0.34}$  for vector boson fusion,  $0.92^{+0.38}_{-0.36}$  for  $WH^0$ ,  $ZH^0$  production, and  $2.90^{+1.08}_{-0.94}$  for  $t\bar{t}H^0$  production.
- 7 AAD 13AK use 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The combined signal strength is based on the  $\gamma\gamma$ ,  $ZZ^* \rightarrow 4\ell$ , and  $WW^* \rightarrow \ell\nu\ell\nu$  channels. The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. Reported statistical error value modified following private communication with the experiment.
- 8 AALTONEN 13L combine all CDF results with 9.45–10.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- 9 ABAZOV 13L combine all D0 results with up to 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- 10 AAD 12AI obtain results based on 4.6–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.8–5.9 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. An excess of events over background with a local significance of 5.9  $\sigma$  is observed at  $m_{H^0} = 126$  GeV. The quoted signal strengths are given for  $m_{H^0} = 126$  GeV. See also AAD 12DA.
- 11 CHATRCHYAN 12N obtain results based on 4.9–5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.1–5.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. An excess of events over background with a local significance of 5.0  $\sigma$  is observed at about  $m_{H^0} = 125$  GeV. The combined signal strength is based on the  $\gamma\gamma$ ,  $ZZ^*$ ,  $WW^*$ ,  $\tau^+\tau^-$ , and  $b\bar{b}$  channels. The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. See also CHATRCHYAN 13Y.

### $WW^*$ Final State

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.08±0.18 -0.16 OUR AVERAGE</b>			
1.09±0.18 -0.16	1,2 AAD	16J LHC	$pp$ , 7, 8 TeV
0.94±0.85 -0.83	3 AALTONEN	13M TEVA	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.22±0.23 -0.21	2 AAD	16J ATLS	$pp$ , 7, 8 TeV
0.90±0.23 -0.21	2 AAD	16J CMS	$pp$ , 7, 8 TeV
1.18±0.16±0.17 -0.14	4 AAD	16K ATLS	$pp$ , 7, 8 TeV
1.09±0.16±0.17 -0.15-0.14	5 AAD	15AA ATLS	$pp$ , 7, 8 TeV
3.0 ± 1.3 ± 1.0 1.1 -0.7	6 AAD	15AQ ATLS	$pp \rightarrow H^0 W/Z X$ , 7, 8 TeV
1.16±0.16±0.18 -0.15-0.15	7 AAD	15AQ ATLS	$pp$ , 7, 8 TeV
0.72±0.12±0.10±0.12 -0.10	8 CHATRCHYAN14G	CMS	$pp$ , 7, 8 TeV
0.99±0.31 -0.28	9 AAD	13AK ATLS	$pp$ , 7 and 8 TeV
0.00±1.78 -0.00	10 AALTONEN	13L CDF	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
1.90±1.63 -1.52	11 ABAZOV	13L D0	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
1.3 ± 0.5	12 AAD	12AI ATLS	$pp \rightarrow H^0 X$ , 7, 8 TeV
0.5 ± 0.6	12 AAD	12AI ATLS	$pp \rightarrow H^0 X$ , 7 TeV
1.9 ± 0.7	12 AAD	12AI ATLS	$pp \rightarrow H^0 X$ , 8 TeV
0.60±0.42 -0.37	13 CHATRCHYAN12N	CMS	$pp \rightarrow H^0 X$ , 7, 8 TeV

- 1 AAD 16J perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and 8 TeV. The signal strengths for individual production processes are  $0.84^{+0.17}_{-0.17}$  for gluon fusion,  $1.2^{+0.4}_{-0.4}$  for vector boson fusion,  $1.6^{+1.2}_{-1.0}$  for  $WH^0$  production,  $5.9^{+2.6}_{-2.2}$  for  $ZH^0$  production, and  $5.0^{+1.8}_{-1.7}$  for  $t\bar{t}H^0$  production.
- 2 In the fit, relative production cross sections are fixed to those in the Standard Model. The quoted signal strength is given for  $m_{H^0} = 125.09$  GeV.
- 3 AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations with up to 10.0 fb<sup>-1</sup> and 9.7 fb<sup>-1</sup>, respectively, of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- 4 AAD 16K use up to 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- 5 AAD 15AA use 4.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The signal strength for the gluon fusion and vector boson fusion mode is  $1.02^{+0.19}_{-0.18} \pm 0.22$  and  $1.27^{+0.44}_{-0.40} \pm 0.30$ , respectively. The quoted signal strengths are given for  $m_{H^0} = 125.36$  GeV.
- 6 AAD 15AQ use 4.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.

See key on page 601

## Gauge &amp; Higgs Boson Particle Listings

 $H^0$ 

- <sup>7</sup> AAD 15AQ combine their result on  $W/ZH^0$  production with the results of AAD 15AA (gluon fusion and vector boson fusion, slightly updated). The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- <sup>8</sup> CHATRCHYAN 14G use  $4.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and  $19.4 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The last uncertainty in the measurement is theory systematics. The quoted signal strength is given for  $m_{H^0} = 125.6$  GeV.
- <sup>9</sup> AAD 13AK use  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and  $20.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. Superseded by AAD 15AA.
- <sup>10</sup> AALTONEN 13L combine all CDF results with  $9.45\text{--}10.0 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>11</sup> ABAZOV 13L combine all D0 results with up to  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>12</sup> AAD 12AI obtain results based on  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and  $5.8 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The quoted signal strengths are given for  $m_{H^0} = 126$  GeV. See also AAD 12DA.
- <sup>13</sup> CHATRCHYAN 12N obtain results based on  $4.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and  $5.1 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. See also CHATRCHYAN 13Y.

 **$ZZ^*$  Final State**

VALUE	DOCUMENT ID	TECN	COMMENT
$1.29^{+0.26}_{-0.23}$	1,2 AAD	16J LHC	$pp, 7, 8 \text{ TeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.52^{+0.40}_{-0.34}$	<sup>2</sup> AAD	16J ATLS	$pp, 7, 8 \text{ TeV}$
$1.04^{+0.32}_{-0.26}$	<sup>2</sup> AAD	16J CMS	$pp, 7, 8 \text{ TeV}$
$1.46^{+0.35+0.19}_{-0.31-0.13}$	<sup>3</sup> AAD	16K ATLS	$pp, 7, 8 \text{ TeV}$
$1.44^{+0.34+0.21}_{-0.31-0.11}$	<sup>4</sup> AAD	15F ATLS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$
	<sup>5</sup> AAD	14AR ATLS	$pp, 8 \text{ TeV, differential cross section}$
$0.93^{+0.26+0.13}_{-0.23-0.09}$	<sup>6</sup> CHATRCHYAN14AA	CMS	$pp, 7, 8 \text{ TeV}$
$1.43^{+0.40}_{-0.35}$	<sup>7</sup> AAD	13AK ATLS	$pp, 7 \text{ and } 8 \text{ TeV}$
$0.80^{+0.35}_{-0.28}$	<sup>8</sup> CHATRCHYAN13J	CMS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$
$1.2 \pm 0.6$	<sup>9</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$
$1.4 \pm 1.1$	<sup>9</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 7 \text{ TeV}$
$1.1 \pm 0.8$	<sup>9</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 8 \text{ TeV}$
$0.73^{+0.45}_{-0.33}$	<sup>10</sup> CHATRCHYAN12N	CMS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$

- <sup>1</sup> AAD 16J perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and  $8 \text{ TeV}$ . The signal strengths for individual production processes are  $1.13^{+0.23}_{-0.31}$  for gluon fusion and  $0.1^{+1.1}_{-0.6}$  for vector boson fusion.
- <sup>2</sup> In the fit, relative production cross sections are fixed to those in the Standard Model. The quoted signal strength is given for  $m_{H^0} = 125.09$  GeV.
- <sup>3</sup> AAD 16K use up to  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and up to  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- <sup>4</sup> AAD 15F use  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV. The signal strength for the gluon fusion production mode is  $1.66^{+0.45+0.25}_{-0.41-0.15}$ , while the signal strength for the vector boson fusion production mode is  $0.26^{+1.60+0.36}_{-0.91-0.23}$ .
- <sup>5</sup> AAD 14AR measure the cross section for  $pp \rightarrow H^0 X, H^0 \rightarrow ZZ^*$  using  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . They give  $\sigma \cdot B = 2.11^{+0.53}_{-0.47} \pm 0.08 \text{ fb}$  in their fiducial region, where  $1.30 \pm 0.13 \text{ fb}$  is expected in the Standard Model for  $m_{H^0} = 125.4$  GeV. Various differential cross sections are also given, which are in agreement with the Standard Model expectations.
- <sup>6</sup> CHATRCHYAN 14AA use  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.6$  GeV. The signal strength for the gluon fusion and  $t\bar{t}H$  production mode is  $0.80^{+0.46}_{-0.36}$ , while the signal strength for the vector boson fusion and  $WH^0, ZH^0$  production mode is  $1.7^{+2.2}_{-2.1}$ .
- <sup>7</sup> AAD 13AK use  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV.
- <sup>8</sup> CHATRCHYAN 13J obtain results based on  $ZZ \rightarrow 4\ell$  final states in  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $12.2 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.8$  GeV. Superseded by CHATRCHYAN 14AA.
- <sup>9</sup> AAD 12AI obtain results based on  $4.7\text{--}4.8 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $5.8 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strengths are given for  $m_{H^0} = 126$  GeV. See also AAD 12DA.
- <sup>10</sup> CHATRCHYAN 12N obtain results based on  $4.9\text{--}5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $5.1\text{--}5.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . An excess of events over background with a local significance of  $5.0 \sigma$  is observed at about  $m_{H^0} = 125$  GeV. The quoted signal strengths are given for  $m_{H^0} = 125.5$  GeV. See also CHATRCHYAN 12BY and CHATRCHYAN 13Y.

 **$\gamma\gamma$  Final State**

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.16 \pm 0.18</math> OUR AVERAGE</b>			
$1.14^{+0.19}_{-0.18}$	<sup>1,2</sup> AAD	16J LHC	$pp, 7, 8 \text{ TeV}$
$5.97^{+3.39}_{-3.12}$	<sup>3</sup> AALTONEN	13M TEVA	$p\bar{p} \rightarrow H^0 X, 1.96 \text{ TeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.14^{+0.27}_{-0.25}$	<sup>2</sup> AAD	16J ATLS	$pp, 7, 8 \text{ TeV}$
$1.11^{+0.25}_{-0.23}$	<sup>2</sup> AAD	16J CMS	$pp, 7, 8 \text{ TeV}$
	<sup>4</sup> KHACHATRY...16B	CMS	$H^0 \rightarrow \gamma^* \gamma \rightarrow \ell^+ \ell^- \gamma$
	<sup>5</sup> KHACHATRY...16G	CMS	differential cross section
$1.17 \pm 0.23^{+0.10+0.12}_{-0.08-0.08}$	<sup>6</sup> AAD	14Bc ATLS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$
	<sup>7</sup> AAD	14Bj ATLS	$pp, 8 \text{ TeV, differential cross section}$
	<sup>8</sup> KHACHATRY...14P	CMS	$pp, 7, 8 \text{ TeV}$
	<sup>9</sup> AAD	13AK ATLS	$pp, 7 \text{ and } 8 \text{ TeV}$
$1.14 \pm 0.21^{+0.09+0.13}_{-0.05-0.09}$	<sup>10</sup> AALTONEN	13L CDF	$p\bar{p} \rightarrow H^0 X, 1.96 \text{ TeV}$
$1.55^{+0.33}_{-0.28}$	<sup>11</sup> ABAZOV	13L D0	$p\bar{p} \rightarrow H^0 X, 1.96 \text{ TeV}$
$7.81^{+4.61}_{-4.42}$	<sup>12</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$
$4.20^{+4.60}_{-4.20}$	<sup>12</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 7 \text{ TeV}$
$1.8 \pm 0.5$	<sup>12</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 8 \text{ TeV}$
$2.2 \pm 0.7$	<sup>12</sup> AAD	12AI ATLS	$pp \rightarrow H^0 X, 8 \text{ TeV}$
$1.5 \pm 0.6$	<sup>13</sup> CHATRCHYAN12N	CMS	$pp \rightarrow H^0 X, 7, 8 \text{ TeV}$
$1.54^{+0.46}_{-0.42}$			

- <sup>1</sup> AAD 16J perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and  $8 \text{ TeV}$ . The signal strengths for individual production processes are  $1.10^{+0.23}_{-0.22}$  for gluon fusion,  $1.3^{+0.5}_{-0.5}$  for vector boson fusion,  $0.5^{+1.3}_{-1.2}$  for  $WH^0$  production,  $0.5^{+3.0}_{-2.5}$  for  $ZH^0$  production, and  $2.2^{+1.6}_{-1.3}$  for  $t\bar{t}H^0$  production.
- <sup>2</sup> In the fit, relative production cross sections are fixed to those in the Standard Model. The quoted signal strength is given for  $m_{H^0} = 125.09$  GeV.
- <sup>3</sup> AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations with up to  $10.0 \text{ fb}^{-1}$  and  $9.7 \text{ fb}^{-1}$ , respectively, of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>4</sup> KHACHATRYAN 16B search for  $H^0 \rightarrow \gamma^* \gamma \rightarrow e^+ e^- \gamma$  and  $\mu^+ \mu^- \gamma$  (with  $m(\ell^+ \ell^-) < 20 \text{ GeV}$ ) in  $19.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8 \text{ TeV}$ . An upper limit of 6.7 times the Standard Model expectation is obtained at 95% CL. See their Fig. 6 for limits on individual channels.
- <sup>5</sup> KHACHATRYAN 16G measure fiducial and differential cross sections of the process  $pp \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$  at  $E_{\text{cm}} = 8 \text{ TeV}$  with  $19.7 \text{ fb}^{-1}$ . See their Figs. 4–6 and Table 1 for data.
- <sup>6</sup> AAD 14Bc use  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The last uncertainty in the measurement is theory systematics. The quoted signal strength is given for  $m_{H^0} = 125.4$  GeV. The signal strengths for the individual production modes are:  $1.32 \pm 0.38$  for gluon fusion,  $0.8 \pm 0.7$  for vector boson fusion,  $1.0 \pm 1.6$  for  $WH^0$  production,  $0.1^{+3.7}_{-0.1}$  for  $ZH^0$  production, and  $1.6^{+2.7}_{-1.8}$  for  $t\bar{t}H^0$  production.
- <sup>7</sup> AAD 14Bj measure fiducial and differential cross sections of the process  $pp \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$  at  $E_{\text{cm}} = 8 \text{ TeV}$  with  $20.3 \text{ fb}^{-1}$ . See their Table 3 and Figs. 3–12 for data.
- <sup>8</sup> KHACHATRYAN 14P use  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The last uncertainty in the measurement is theory systematics. The quoted signal strength is given for  $m_{H^0} = 124.7$  GeV. The signal strength for the gluon fusion and  $t\bar{t}H$  production mode is  $1.13^{+0.37}_{-0.31}$ , while the signal strength for the vector boson fusion and  $WH^0, ZH^0$  production mode is  $1.16^{+0.63}_{-0.58}$ .
- <sup>9</sup> AAD 13AK use  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV.
- <sup>10</sup> AALTONEN 13L combine all CDF results with  $9.45\text{--}10.0 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>11</sup> ABAZOV 13L combine all D0 results with up to  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>12</sup> AAD 12AI obtain results based on  $4.8 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $5.9 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strengths are given for  $m_{H^0} = 126$  GeV. See also AAD 12DA.
- <sup>13</sup> CHATRCHYAN 12N obtain results based on  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $5.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. See also CHATRCHYAN 13Y.

 **$b\bar{b}$  Final State**

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.82 \pm 0.30</math> OUR AVERAGE</b>			Error includes scale factor of 1.1.
$0.70^{+0.29}_{-0.27}$	<sup>1,2</sup> AAD	16J LHC	$pp, 7, 8 \text{ TeV}$
$1.59^{+0.69}_{-0.72}$	<sup>3</sup> AALTONEN	13M TEVA	$p\bar{p} \rightarrow H^0 X, 1.96 \text{ TeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.62 \pm 0.37$	<sup>2</sup> AAD	16J ATLS	$pp, 7, 8 \text{ TeV}$
$0.81^{+0.45}_{-0.43}$	<sup>2</sup> AAD	16J CMS	$pp, 7, 8 \text{ TeV}$
$0.63^{+0.31+0.24}_{-0.30-0.23}$	<sup>4</sup> AAD	16K ATLS	$pp, 7, 8 \text{ TeV}$
$0.52 \pm 0.32 \pm 0.24$	<sup>5</sup> AAD	16G ATLS	$pp \rightarrow H^0 W/ZX, 7, 8 \text{ TeV}$
$2.8^{+1.6}_{-1.4}$	<sup>6</sup> KHACHATRY...15z	CMS	$pp \rightarrow H^0 X, \text{VBF}, 8 \text{ TeV}$
$1.03^{+0.44}_{-0.42}$	<sup>7</sup> KHACHATRY...15z	CMS	$pp, 8 \text{ TeV, combined}$
$1.0 \pm 0.5$	<sup>8</sup> CHATRCHYAN14AI	CMS	$pp \rightarrow H^0 W/ZX, 7, 8 \text{ TeV}$
$1.72^{+0.92}_{-0.87}$	<sup>9</sup> AALTONEN	13L CDF	$p\bar{p} \rightarrow H^0 X, 1.96 \text{ TeV}$
$1.23^{+1.24}_{-1.17}$	<sup>10</sup> ABAZOV	13L D0	$p\bar{p} \rightarrow H^0 X, 1.96 \text{ TeV}$
$0.5 \pm 0.2$	<sup>11</sup> AAD	12AI ATLS	$pp \rightarrow H^0 W/ZX, 7 \text{ TeV}$
	<sup>12</sup> AALTONEN	12T TEVA	$p\bar{p} \rightarrow H^0 W/ZX, 1.96 \text{ TeV}$
$0.48^{+0.81}_{-0.70}$	<sup>13</sup> CHATRCHYAN12N	CMS	$pp \rightarrow H^0 W/ZX, 7, 8 \text{ TeV}$

# Gauge & Higgs Boson Particle Listings

## $H^0$

- <sup>1</sup> AAD 16j perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and 8 TeV. The signal strengths for individual production processes are  $1.0^{+0.5}_{-0.5}$  for  $WH^0$  production,  $0.4^{+0.4}_{-0.4}$  for  $ZH^0$  production, and  $1.1^{+1.0}_{-1.0}$  for  $t\bar{t}H^0$  production.
- <sup>2</sup> In the fit, relative production cross sections are fixed to those in the Standard Model. The quoted signal strength is given for  $m_{H^0} = 125.09$  GeV.
- <sup>3</sup> AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations with up to 10.0 fb<sup>-1</sup> and 9.7 fb<sup>-1</sup>, respectively, of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>4</sup> AAD 16k use up to 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- <sup>5</sup> AAD 15g use 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- <sup>6</sup> KHACHATRYAN 15z search for vector-boson fusion production of  $H^0$  decaying to  $b\bar{b}$  in up to 19.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>7</sup> KHACHATRYAN 15z combined vector boson fusion,  $WH^0$ ,  $ZH^0$  production, and  $t\bar{t}H^0$  production results. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>8</sup> CHATRCHYAN 14Ai use up to 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 18.9 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV. See also CHATRCHYAN 14Aj.
- <sup>9</sup> AALTONEN 13L combine all CDF results with 9.45–10.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>10</sup> ABAZOV 13L combine all D0 results with up to 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>11</sup> AAD 12Ai obtain results based on 4.6–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. The quoted signal strengths are given in their Fig. 10 for  $m_{H^0} = 126$  GeV. See also Fig. 13 of AAD 12Da.
- <sup>12</sup> AALTONEN 12T combine AALTONEN 12Q, AALTONEN 12R, AALTONEN 12S, ABAZOV 12P, and ABAZOV 12K. An excess of events over background is observed which is most significant in the region  $m_{H^0} = 120$ –135 GeV, with a local significance of up to 3.3  $\sigma$ . The local significance at  $m_{H^0} = 125$  GeV is 2.8  $\sigma$ , which corresponds to  $(\sigma(H^0 W) + \sigma(H^0 Z)) \cdot \text{B}(H^0 \rightarrow b\bar{b}) = (0.23^{+0.09}_{-0.08})$  pb, compared to the Standard Model expectation at  $m_{H^0} = 125$  GeV of  $0.12 \pm 0.01$  pb. Superseded by AALTONEN 13M.
- <sup>13</sup> CHATRCHYAN 12N obtain results based on 5.0 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.1 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. See also CHATRCHYAN 13Y.

### $\mu^+ \mu^-$ Final State

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<7.4	95	<sup>1</sup> KHACHATRY..15H	CMS	$pp \rightarrow H^0 X$ , 7, 8 TeV
<7.0	95	<sup>2</sup> AAD	14As ATLS	$pp \rightarrow H^0 X$ , 7, 8 TeV
<sup>1</sup> KHACHATRYAN 15H use 5.0 fb <sup>-1</sup> of $pp$ collisions at $E_{\text{cm}} = 7$ TeV and 19.7 fb <sup>-1</sup> at 8 TeV. The quoted signal strength is given for $m_{H^0} = 125$ GeV.				
<sup>2</sup> AAD 14As search for $H^0 \rightarrow \mu^+ \mu^-$ in 4.5 fb <sup>-1</sup> of $pp$ collisions at $E_{\text{cm}} = 7$ TeV and 20.3 fb <sup>-1</sup> at $E_{\text{cm}} = 8$ TeV. The quoted signal strength is given for $m_{H^0} = 125.5$ GeV.				

### $\tau^+ \tau^-$ Final State

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.12±0.23 OUR AVERAGE</b>			
$1.11^{+0.24}_{-0.22}$	<sup>1,2</sup> AAD	16j LHC	$pp$ , 7, 8 TeV
$1.68^{+2.28}_{-1.68}$	<sup>3</sup> AALTONEN	13M TEVA	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.41^{+0.40}_{-0.36}$	<sup>2</sup> AAD	16j ATLS	$pp$ , 7, 8 TeV
$0.88^{+0.30}_{-0.28}$	<sup>2</sup> AAD	16j CMS	$pp$ , 7, 8 TeV
$1.44^{+0.30+0.29}_{-0.29-0.23}$	<sup>4</sup> AAD	16k ATLS	$pp$ , 7, 8 TeV
$1.43^{+0.27+0.32}_{-0.26-0.25} \pm 0.09$	<sup>5</sup> AAD	15AH ATLS	$pp \rightarrow H^0 X$ , 7, 8 TeV
$0.78 \pm 0.27$	<sup>6</sup> CHATRCHYAN 14K	CMS	$pp \rightarrow H^0 X$ , 7, 8 TeV
$0.00^{+8.44}_{-0.00}$	<sup>7</sup> AALTONEN	13L CDF	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
$3.96^{+4.11}_{-3.38}$	<sup>8</sup> ABAZOV	13L D0	$p\bar{p} \rightarrow H^0 X$ , 1.96 TeV
$0.4^{+1.6}_{-2.0}$	<sup>9</sup> AAD	12Ai ATLS	$pp \rightarrow H^0 X$ , 7 TeV
$0.09^{+0.76}_{-0.74}$	<sup>10</sup> CHATRCHYAN 12N	CMS	$pp \rightarrow H^0 X$ , 7, 8 TeV

- <sup>1</sup> AAD 16j perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and 8 TeV. The signal strengths for individual production processes are  $1.0^{+0.6}_{-0.6}$  for gluon fusion,  $1.3^{+0.4}_{-0.4}$  for vector boson fusion,  $-1.4^{+1.4}_{-1.4}$  for  $WH^0$  production,  $2.2^{+2.2}_{-1.8}$  for  $ZH^0$  production, and  $-1.9^{+3.7}_{-3.3}$  for  $t\bar{t}H^0$  production.
- <sup>2</sup> In the fit, relative production cross sections are fixed to those in the Standard Model. The quoted signal strength is given for  $m_{H^0} = 125.09$  GeV.
- <sup>3</sup> AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations with up to 10.0 fb<sup>-1</sup> and 9.7 fb<sup>-1</sup>, respectively, of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>4</sup> AAD 16k use up to 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.

- <sup>5</sup> AAD 15AH use 4.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The third uncertainty in the measurement is theory systematics. The signal strength for the gluon fusion mode is  $2.0 \pm 0.8^{+1.2}_{-0.8} \pm 0.3$  and that for vector boson fusion and  $W/ZH^0$  production modes is  $1.24^{+0.49+0.31}_{-0.45-0.29} \pm 0.08$ . The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- <sup>6</sup> CHATRCHYAN 14K use 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV. See also CHATRCHYAN 14AJ.
- <sup>7</sup> AALTONEN 13L combine all CDF results with 9.45–10.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>8</sup> ABAZOV 13L combine all D0 results with up to 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>9</sup> AAD 12Ai obtain results based on 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. The quoted signal strengths are given in their Fig. 10 for  $m_{H^0} = 126$  GeV. See also Fig. 13 of AAD 12DA.
- <sup>10</sup> CHATRCHYAN 12N obtain results based on 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.1 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.5$  GeV. See also CHATRCHYAN 13Y.

### $Z\gamma$ Final State

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<11	95	<sup>1</sup> AAD	14j ATLS	$pp \rightarrow H^0 X$ , 7, 8 TeV
< 9.5	95	<sup>2</sup> CHATRCHYAN 13BK	CMS	$pp \rightarrow H^0 X$ , 7, 8 TeV
<sup>1</sup> AAD 14j search for $H^0 \rightarrow Z\gamma \rightarrow \ell\ell\gamma$ in 4.5 fb <sup>-1</sup> of $pp$ collisions at $E_{\text{cm}} = 7$ TeV and 20.3 fb <sup>-1</sup> at $E_{\text{cm}} = 8$ TeV. The quoted signal strength is given for $m_{H^0} = 125.5$ GeV.				
<sup>2</sup> CHATRCHYAN 13BK search for $H^0 \rightarrow Z\gamma \rightarrow \ell\ell\gamma$ in 5.0 fb <sup>-1</sup> of $pp$ collisions at $E_{\text{cm}} = 7$ TeV and 19.6 fb <sup>-1</sup> at $E_{\text{cm}} = 8$ TeV. A limit on cross section times branching ratio which corresponds to (4–25) times the expected Standard Model cross section is given in the range $m_{H^0} = 120$ –160 GeV at 95% CL. The quoted limit is given for $m_{H^0} = 125$ GeV, where 10 is expected for no signal.				

### $t\bar{t}H^0$ Production

Signal strength relative to the Standard Model cross section.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.3 <math>\pm 0.7</math> –0.6</b>		<sup>1,2</sup> AAD	16j LHC	$pp$ , 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.9^{+0.8}_{-0.7}$		<sup>2</sup> AAD	16j ATLS	$pp$ , 7, 8 TeV
$2.9^{+1.0}_{-0.9}$		<sup>2</sup> AAD	16j CMS	$pp$ , 7, 8 TeV
$1.81^{+0.52+0.58+0.31}_{-0.50-0.55-0.12}$		<sup>3</sup> AAD	16k ATLS	$pp$ , 7, 8 TeV
$1.4^{+2.1+0.6}_{-1.4-0.3}$		<sup>4</sup> AAD	15 ATLS	$pp$ , 7, 8 TeV
$1.5 \pm 1.1$		<sup>5</sup> AAD	15bc ATLS	$pp$ , 8 TeV
$2.1^{+1.4}_{-1.2}$		<sup>6</sup> AAD	15T ATLS	$pp$ , 8 TeV
$1.2^{+1.6}_{-1.5}$		<sup>7</sup> KHACHATRY..15AN	CMS	$pp$ , 8 TeV
$2.8^{+1.0}_{-0.9}$		<sup>8</sup> KHACHATRY..14H	CMS	$pp$ , 7, 8 TeV
$9.49^{+6.60}_{-6.28}$		<sup>9</sup> AALTONEN	13L CDF	$p\bar{p}$ , 1.96 TeV
<5.8	95	<sup>10</sup> CHATRCHYAN 13X	CMS	$pp \rightarrow H^0 t\bar{t}X$

- <sup>1</sup> AAD 16j perform fits to the ATLAS and CMS data at  $E_{\text{cm}} = 7$  and 8 TeV.
- <sup>2</sup> In the fit, relative branching ratios are fixed to those in the Standard Model. The quoted signal strength is given for  $m_{H^0} = 125.09$  GeV.
- <sup>3</sup> AAD 16k use up to 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The third uncertainty in the measurement is theory systematics. The quoted signal strength is given for  $m_{H^0} = 125.36$  GeV.
- <sup>4</sup> AAD 15 search for  $t\bar{t}H^0$  production with  $H^0$  decaying to  $\gamma\gamma$  in 4.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted result on the signal strength is equivalent to an upper limit of 6.7 at 95% CL and is given for  $m_{H^0} = 125.4$  GeV.
- <sup>5</sup> AAD 15bc search for  $t\bar{t}H^0$  production with  $H^0$  decaying to  $b\bar{b}$  in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. The corresponding upper limit is 3.4 at 95% CL. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>6</sup> AAD 15T search for  $t\bar{t}H^0$  production with  $H^0$  resulting in multilepton final states (mainly from  $WW^*$ ,  $\tau\tau$ ,  $ZZ^*$ ) in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. The quoted result on the signal strength is given for  $m_{H^0} = 125$  GeV and corresponds to an upper limit of 4.7 at 95% CL. The data sample is independent from AAD 15 and AAD 15bc.
- <sup>7</sup> KHACHATRYAN 15AN search for  $t\bar{t}H^0$  production with  $H^0$  decaying to  $b\bar{b}$  in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. The quoted result on the signal strength is equivalent to an upper limit of 4.2 at 95% CL and is given for  $m_{H^0} = 125$  GeV.
- <sup>8</sup> KHACHATRYAN 14H search for  $t\bar{t}H^0$  production with  $H^0$  decaying to  $b\bar{b}$ ,  $\tau\tau$ ,  $\gamma\gamma$ ,  $WW^*$ , and  $ZZ^*$ , in 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. The quoted signal strength is given for  $m_{H^0} = 125.6$  GeV.
- <sup>9</sup> AALTONEN 13L combine all CDF results with 9.45–10.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The quoted signal strength is given for  $m_{H^0} = 125$  GeV.
- <sup>10</sup> CHATRCHYAN 13X search for  $t\bar{t}H^0$  production followed by  $H^0 \rightarrow b\bar{b}$ , one top decaying to  $\ell\nu$  and the other to either  $\ell\nu$  or  $q\bar{q}$  in 5.0 fb<sup>-1</sup> and 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  and 8 TeV. A limit on cross section times branching ratio which corresponds to (4.0–8.6) times the expected Standard Model cross section is given for  $m_{H^0} = 110$ –140 GeV at 95% CL. The quoted limit is given for  $m_{H^0} = 125$  GeV, where 5.2 is expected for no signal.

See key on page 601

# Gauge & Higgs Boson Particle Listings

## $H^0$ , Neutral Higgs Bosons, Searches for

### $H^0$ REFERENCES

AAD	16	PL B753 69	G. Aad et al.	(ATLAS Collab.)
AAD	16J	arXiv:1606.02266	G. Aad et al.	(ATLAS and CMS Collabs.)
JHEP to be published				
AAD	16K	EPJ C76 6	G. Aad et al.	(ATLAS Collab.)
KHACHATRYAN...	16B	PL B753 341	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	16G	EPJ C76 13	V. Khachatryan et al.	(CMS Collab.)
AAD	15	PL B740 222	G. Aad et al.	(ATLAS Collab.)
AAD	15AA	PR D92 012006	G. Aad et al.	(ATLAS Collab.)
AAD	15AH	JHEP 1504 117	G. Aad et al.	(ATLAS Collab.)
AAD	15AQ	JHEP 1508 137	G. Aad et al.	(ATLAS Collab.)
AAD	15AX	EPJ C75 231	G. Aad et al.	(ATLAS Collab.)
AAD	15B	PRL 114 191803	G. Aad et al.	(ATLAS and CMS Collabs.)
AAD	15BC	EPJ C75 349	G. Aad et al.	(ATLAS Collab.)
AAD	15BD	EPJ C75 337	G. Aad et al.	(ATLAS Collab.)
AAD	15BE	EPJ C75 335	G. Aad et al.	(ATLAS Collab.)
AAD	15CI	EPJ C75 476	G. Aad et al.	(ATLAS Collab.)
AAD	15F	PR D91 012006	G. Aad et al.	(ATLAS Collab.)
AAD	15G	JHEP 1501 069	G. Aad et al.	(ATLAS Collab.)
AAD	15I	PRL 114 121801	G. Aad et al.	(ATLAS Collab.)
AAD	15P	PRL 115 091801	G. Aad et al.	(ATLAS Collab.)
AAD	15T	PL B749 519	G. Aad et al.	(ATLAS Collab.)
AALTONEN	15	PRL 114 151802	T. Aaltonen et al.	(CDF and DO Collabs.)
AALTONEN	15B	PRL 114 141802	T. Aaltonen et al.	(CDF Collab.)
KHACHATRYAN...	15AM	EPJ C75 212	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15AN	EPJ C75 251	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15BA	PR D92 072010	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15H	PL B744 184	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15Q	PL B749 337	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15Y	PR D92 012004	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15Z	PR D92 032008	V. Khachatryan et al.	(CMS Collab.)
AAD	14AR	PL B738 234	G. Aad et al.	(ATLAS Collab.)
AAD	14AS	PL B738 68	G. Aad et al.	(ATLAS Collab.)
AAD	14BC	PR D90 112015	G. Aad et al.	(ATLAS Collab.)
AAD	14BJ	JHEP 1409 112	G. Aad et al.	(ATLAS Collab.)
AAD	14J	PL B732 8	G. Aad et al.	(ATLAS Collab.)
AAD	14O	PRL 112 201802	G. Aad et al.	(ATLAS Collab.)
AAD	14W	PR D90 052004	G. Aad et al.	(ATLAS Collab.)
ABAZOV	14F	PRL 113 161802	V.M. Abazov et al.	(DO Collab.)
CHATRCHYAN	14AA	PR D89 092007	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14AI	PR D89 012003	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14AJ	NATP 10 557	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14B	EPJ C74 2980	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14G	JHEP 1401 096	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14K	JHEP 1405 104	S. Chatrchyan et al.	(CMS Collab.)
KHACHATRYAN...	14D	PL B736 64	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	14H	JHEP 1409 087	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	14P	EPJ C74 3076	V. Khachatryan et al.	(CMS Collab.)
AAD	13AJ	PL B726 120	G. Aad et al.	(ATLAS Collab.)
AAD	13AK	PL B726 88	G. Aad et al.	(ATLAS Collab.)
Also				
AALTONEN	13L	PR D88 052013	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	13M	PR D88 052014	T. Aaltonen et al.	(CDF and DO Collabs.)
ABAZOV	13L	PR D88 052011	V.M. Abazov et al.	(DO Collab.)
CHATRCHYAN	13BK	PL B726 587	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13J	PRL 110 081803	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13X	JHEP 1305 145	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13Y	JHEP 1306 081	S. Chatrchyan et al.	(CMS Collab.)
HEINEMEYER	13A	arXiv:1307.1347	S. Heinemeyer et al.	(LHC Higgs CS Working Group)
AAD	12AI	PL B716 1	G. Aad et al.	(ATLAS Collab.)
AAD	12DA	SCI 338 1576	G. Aad et al.	(ATLAS Collab.)
AALTONEN	12Q	PRL 109 111803	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12R	PRL 109 111804	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12S	PRL 109 111805	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12T	PRL 109 071804	T. Aaltonen et al.	(CDF and DO Collabs.)
ABAZOV	12K	PL B716 285	V.M. Abazov et al.	(DO Collab.)
ABAZOV	12O	PRL 109 121803	V.M. Abazov et al.	(DO Collab.)
ABAZOV	12P	PRL 109 121804	V.M. Abazov et al.	(DO Collab.)
CHATRCHYAN	12BY	SCI 338 1569	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12N	PL B716 30	S. Chatrchyan et al.	(CMS Collab.)
DITTMAYER	12	arXiv:1201.3084	S. Dittmaier et al.	(LHC Higgs CS Working Group)
DITTMAYER	11	arXiv:1101.0593	S. Dittmaier et al.	(LHC Higgs CS Working Group)

## Neutral Higgs Bosons, Searches for

### CONTENTS:

- Mass Limits for Neutral Higgs Bosons in Supersymmetric Models
  - Mass Limits for  $H_1^0$  (Higgs Boson) in Supersymmetric Models
  - Mass Limits for  $A^0$  (Pseudoscalar Higgs Boson) in Supersymmetric Models
- Mass Limits for Neutral Higgs Bosons in Extended Higgs Models
  - Mass Limits in General two-Higgs-doublet Models
  - Mass Limits for  $H^0$  with Vanishing Yukawa Couplings
  - Mass Limits for  $H^0$  Decaying to Invisible Final States
  - Mass Limits for Light  $A^0$
  - Other Mass Limits
- Searches for a Higgs Boson with Standard Model Couplings
  - Direct Mass Limits for  $H^0$
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### MASS LIMITS FOR NEUTRAL HIGGS BOSONS IN SUPERSYMMETRIC MODELS

The minimal supersymmetric model has two complex doublets of Higgs bosons. The resulting physical states are two scalars [ $H_1^0$  and  $H_2^0$ ], where we define  $m_{H_1^0} < m_{H_2^0}$ , a pseudoscalar ( $A^0$ ), and a charged Higgs pair ( $H^\pm$ ).  $H_1^0$  and  $H_2^0$  are also called  $h$  and  $H$  in the literature. There are two free parameters in the Higgs sector which can be chosen to be  $m_{A^0}$  and  $\tan\beta = v_2/v_1$ , the ratio of vacuum expectation values of the two Higgs doublets. Tree-level Higgs masses are constrained by the model to be  $m_{H_1^0} \leq m_Z$ ,  $m_{H_2^0} \geq m_Z$ ,  $m_{A^0} \geq m_{H_1^0}$ , and  $m_{H^\pm} \geq m_W$ .

However, as described in the review on “Status of Higgs Boson Physics” in this Volume these relations are violated by radiative corrections.

Unless otherwise noted, the experiments in  $e^+e^-$  collisions search for the processes  $e^+e^- \rightarrow H_1^0 Z^0$  in the channels used for the Standard Model Higgs searches and  $e^+e^- \rightarrow H_1^0 A^0$  in the final states  $b\bar{b}b\bar{b}$  and  $b\bar{b}\tau^+\tau^-$ . In  $p\bar{p}$  and  $pp$  collisions the experiments search for a variety of processes, as explicitly specified for each entry. Limits on the  $A^0$  mass arise from these direct searches, as well as from the relations valid in the minimal supersymmetric model between  $m_{A^0}$  and  $m_{H_1^0}$ . As discussed

in the review on “Status of Higgs Boson Physics” in this Volume, these relations depend, via potentially large radiative corrections, on the mass of the  $t$  quark and on the supersymmetric parameters, in particular those of the stop sector. These indirect limits are weaker for larger  $t$  and  $\tilde{t}$  masses. To include the radiative corrections to the Higgs masses, unless otherwise stated, the listed papers use theoretical predictions incorporating two-loop corrections, and the results are given for the  $m_h^{\text{max}}$  benchmark scenario, which gives rise to the most conservative upper bound on the mass of  $H_1^0$  for given values of  $m_{A^0}$  and  $\tan\beta$ , see CARENA 99b, CARENA 03, and CARENA 13.

Limits in the low-mass region of  $H_1^0$ , as well as other by now obsolete limits from different techniques, have been removed from this compilation, and can be found in earlier editions of this Review. Unless otherwise stated, the following results assume no invisible  $H_1^0$  or  $A^0$  decays.

The observed signal at about 125 GeV, see section “ $H^0$ ”, can be interpreted as one of the neutral Higgs bosons of supersymmetric models.

### Mass Limits for $H_1^0$ (Higgs Boson) in Supersymmetric Models

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>89.7		<sup>1</sup> ABDALLAH 08B	DLPH	$E_{\text{cm}} \leq 209$ GeV
>92.8	95	<sup>2</sup> SCHAEEL 06B	LEP	$E_{\text{cm}} \leq 209$ GeV
>84.5	95	<sup>3,4</sup> ABBIENDI 04M	OPAL	$E_{\text{cm}} \leq 209$ GeV
>86.0	95	<sup>3,5</sup> ACHARD 02H	L3	$E_{\text{cm}} \leq 209$ GeV, $\tan\beta > 0.4$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>6</sup> KHACHATRYAN...16A	CMS	$H_{1,2}^0/A^0 \rightarrow \mu^+\mu^-$
		<sup>7</sup> AAD 15CE	ATLS	$H_2^0 \rightarrow H^0 H^0$
		<sup>8</sup> KHACHATRYAN...15AY	CMS	$pp \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>9</sup> AAD 14AW	ATLS	$pp \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau\tau$
		<sup>10</sup> KHACHATRYAN...14M	CMS	$pp \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau\tau$
		<sup>11</sup> AAD 13O	ATLS	$pp \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-, \mu^+\mu^-$
		<sup>12</sup> AAIJ 13T	LHCb	$pp \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>13</sup> CHATRCHYAN13AG	CMS	$pp \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>14</sup> AALTONEN 12AQ	TEVA	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>15</sup> AALTONEN 12X	CDF	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>16</sup> ABAZOV 12G	D0	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>17</sup> CHATRCHYAN12K	CMS	$pp \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>18</sup> ABAZOV 11K	D0	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>19</sup> ABAZOV 11W	D0	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>20</sup> AALTONEN 09AR	CDF	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>21</sup> ABBIENDI 03G	OPAL	$H_1^0 \rightarrow A^0 A^0$
>89.8	95	<sup>3,22</sup> HEISTER 02	ALEP	$E_{\text{cm}} \leq 209$ GeV, $\tan\beta > 0.5$

<sup>1</sup> ABDALLAH 08B give limits in eight  $CP$ -conserving benchmark scenarios and some  $CP$ -violating scenarios. See paper for excluded regions for each scenario. Supersedes ABDALLAH 04.

<sup>2</sup> SCHAEEL 06B make a combined analysis of the LEP data. The quoted limit is for the  $m_h^{\text{max}}$  scenario with  $m_t = 174.3$  GeV. In the  $CP$ -violating CPX scenario no lower bound on  $m_{H_1^0}$  can be set at 95% CL. See paper for excluded regions in various scenarios. See

Figs. 2–6 and Tabs. 14–21 for limits on  $\sigma(ZH^0) \cdot B(H^0 \rightarrow b\bar{b}, \tau^+\tau^-)$  and  $\sigma(H_1^0 H_2^0) \cdot B(H_1^0 H_2^0 \rightarrow b\bar{b}, \tau^+\tau^-)$ .

# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

- <sup>3</sup> Search for  $e^+e^- \rightarrow H_1^0 A^0$  in the final states  $b\bar{b}b\bar{b}$  and  $b\bar{b}\tau^+\tau^-$ , and  $e^+e^- \rightarrow H_1^0 Z$ . Universal scalar mass of 1 TeV, SU(2) gaugino mass of 200 GeV, and  $\mu = -200$  GeV are assumed, and two-loop radiative corrections incorporated. The limits hold for  $m_t = 175$  GeV, and for the  $m_h^{\max}$  scenario.
- <sup>4</sup> ABBIENDI 04M exclude  $0.7 < \tan\beta < 1.9$ , assuming  $m_t = 174.3$  GeV. Limits for other MSSM benchmark scenarios, as well as for CP violating cases, are also given.
- <sup>5</sup> ACHARD 02H also search for the final state  $H_1^0 Z \rightarrow 2A^0 q\bar{q}, A^0 \rightarrow q\bar{q}$ . In addition, the MSSM parameter set in the “large- $\mu$ ” and “no-mixing” scenarios are examined.
- <sup>6</sup> KHACHATRYAN 16A search for production of a Higgs boson in gluon fusion and in association with a  $b\bar{b}$  pair followed by the decay  $H_{1,2}^0/A^0 \rightarrow \mu^+\mu^-$  in 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. See their Fig. 7 for the excluded region in the MSSM parameter space in the  $m_h^{\text{mod+}}$  benchmark scenario and Fig. 9 for limits on cross section times branching ratio.
- <sup>7</sup> AAD 15CE search for production of  $H_2^0$  decaying to  $H^0 H^0$  in the final states  $b\bar{b}\tau^+\tau^-$  and  $\gamma\gamma WW^*$  in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV and combine with data from AAD 15H ( $\gamma\gamma b\bar{b}$ ) and AAD 15BK ( $b\bar{b}b\bar{b}$ ). See their Fig. 7 for excluded regions in the parameter space in several scenarios.
- <sup>8</sup> KHACHATRYAN 15AY search for production of a Higgs boson in association with a  $b$  quark in the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$  in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV and combine with CHATRCHYAN 13AG 7 TeV data. See their Fig. 6 for the limits on cross section times branching ratio for  $m_{A^0} = 100\text{--}900$  GeV and Figs. 7–9 for the excluded region in the MSSM parameter space in various benchmark scenarios.
- <sup>9</sup> AAD 14AW search for production of a Higgs boson followed by the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  in 19.5–20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 11 for the limits on cross section times branching ratio and their Figs. 9 and 10 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 140$  GeV, the region  $\tan\beta > 5.4$  is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>10</sup> KHACHATRYAN 14M search for production of a Higgs boson in gluon fusion and in association with a  $b$  quark followed by the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  in 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. See their Figs. 7 and 8 for one- and two-dimensional limits on cross section times branching ratio and their Figs. 5 and 6 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 140$  GeV, the region  $\tan\beta > 3.8$  is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>11</sup> AAD 13o search for production of a Higgs boson in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  and  $\mu^+\mu^-$  with 4.7–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 6 for the excluded region in the MSSM parameter space and their Fig. 7 for the limits on cross section times branching ratio. For  $m_{A^0} = 110\text{--}170$  GeV,  $\tan\beta \gtrsim 10$  is excluded, and for  $\tan\beta = 50$ ,  $m_{A^0}$  below 470 GeV is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>12</sup> AAIJ 13T search for production of a Higgs boson in the forward region in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  in 1.0 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 2 for the limits on cross section times branching ratio and the excluded region in the MSSM parameter space.
- <sup>13</sup> CHATRCHYAN 13AG search for production of a Higgs boson in association with a  $b$  quark in the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$  in 2.7–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 6 for the excluded region in the MSSM parameter space and Fig. 5 for the limits on cross section times branching ratio. For  $m_{A^0} = 90\text{--}350$  GeV, upper bounds on  $\tan\beta$  of 18–42 at 95% CL are obtained in the  $m_h^{\max}$  scenario with  $\mu = +200$  GeV.
- <sup>14</sup> AALTONEN 12AQ combine AALTONEN 12X and ABAZOV 11K. See their Table I and Fig. 1 for the limit on cross section times branching ratio and Fig. 2 for the excluded region in the MSSM parameter space.
- <sup>15</sup> AALTONEN 12X search for associated production of a Higgs boson and a  $b$  quark in the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$ , with 2.6 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. See their Table III and Fig. 15 for the limit on cross section times branching ratio and Figs. 17, 18 for the excluded region in the MSSM parameter space.
- <sup>16</sup> ABAZOV 12C search for production of a Higgs boson in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  with 7.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV and combine with ABAZOV 11W and ABAZOV 11K. See their Figs. 4, 5, and 6 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 90\text{--}180$  GeV,  $\tan\beta \gtrsim 30$  is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>17</sup> CHATRCHYAN 12K search for production of a Higgs boson in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  with 4.6 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 3 and Table 4 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 160$  GeV, the region  $\tan\beta > 7.1$  is excluded at 95% CL in the  $m_h^{\max}$  scenario. Superseded by KHACHATRYAN 14M.
- <sup>18</sup> ABAZOV 11K search for associated production of a Higgs boson and a  $b$  quark, followed by the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$ , in 5.2 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. See their Fig. 5/Table 2 for the limit on cross section times branching ratio and Fig. 6 for the excluded region in the MSSM parameter space for  $\mu = -200$  GeV.
- <sup>19</sup> ABAZOV 11W search for associated production of a Higgs boson and a  $b$  quark, followed by the decay  $H_{1,2}^0/A^0 \rightarrow \tau\tau$ , in 7.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. See their Fig. 2 for the limit on cross section times branching ratio and for the excluded region in the MSSM parameter space.
- <sup>20</sup> AALTONEN 09AR search for Higgs bosons decaying to  $\tau^+\tau^-$  in two doublet models in 1.8 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. See their Fig. 2 for the limit on  $\sigma \cdot B(H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-)$  for different Higgs masses, and see their Fig. 3 for the excluded region in the MSSM parameter space.
- <sup>21</sup> ABBIENDI 03C search for  $e^+e^- \rightarrow H_1^0 Z$  followed by  $H_1^0 \rightarrow A^0 A^0, A^0 \rightarrow c\bar{c}, gg$ , or  $\tau^+\tau^-$ . In the no-mixing scenario, the region  $m_{H_1^0} = 45\text{--}85$  GeV and  $m_{A^0} = 2\text{--}9.5$  GeV is excluded at 95% CL.
- <sup>22</sup> HEISTER 02 excludes the range  $0.7 < \tan\beta < 2.3$ . A wider range is excluded with different stop mixing assumptions. Updates BARATE 01c.

## Mass Limits for $A^0$ (Pseudoscalar Higgs Boson) in Supersymmetric Models

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>90.4		<sup>1</sup> ABDALLAH 08B	08B DLPH	$E_{\text{cm}} \leq 209$ GeV
>93.4	95	<sup>2</sup> SCHAEEL 06B	06B LEP	$E_{\text{cm}} \leq 209$ GeV
>85.0	95	<sup>3,4</sup> ABBIENDI 04M	04M OPAL	$E_{\text{cm}} \leq 209$ GeV
>86.5	95	<sup>3,5</sup> ACHARD 02H	L3	$E_{\text{cm}} \leq 209$ GeV, $\tan\beta > 0.4$
>90.1	95	<sup>3,6</sup> HEISTER 02	ALEP	$E_{\text{cm}} \leq 209$ GeV, $\tan\beta > 0.5$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>7</sup> KHACHATRYAN...16A	CMS	$H_{1,2}^0/A^0 \rightarrow \mu^+\mu^-$
		<sup>8</sup> KHACHATRYAN...15AY	CMS	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>9</sup> AAD	14AWATLS	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau\tau$
		<sup>10</sup> KHACHATRYAN...14M	CMS	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau\tau$
		<sup>11</sup> AAD	13o ATLS	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-, \mu^+\mu^-$
		<sup>12</sup> AAIJ	13T LHCB	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>13</sup> CHATRCHYAN 13AG	CMS	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>14</sup> AALTONEN	12AQ TEVA	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>15</sup> AALTONEN	12X CDF	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>16</sup> ABAZOV	12G D0	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>17</sup> CHATRCHYAN 12K	CMS	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>18</sup> ABAZOV	11K D0	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow b\bar{b}$
		<sup>19</sup> ABAZOV	11W D0	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + b + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>20</sup> AALTONEN	09AR CDF	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$ , $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$
		<sup>21</sup> ACOSTA	05Q CDF	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X$
		<sup>22</sup> ABBIENDI	03G OPAL	$H_1^0 \rightarrow A^0 A^0$
		<sup>23</sup> AKEROYD	02 RVUE	

<sup>1</sup> ABDALLAH 08B give limits in eight CP-conserving benchmark scenarios and some CP-violating scenarios. See paper for excluded regions for each scenario. Supersedes ABDALLAH 04.

<sup>2</sup> SCHAEEL 06B make a combined analysis of the LEP data. The quoted limit is for the  $m_h^{\max}$  scenario with  $m_t = 174.3$  GeV. In the CP-violating CPX scenario no lower bound on  $m_{H_1^0}$  can be set at 95% CL. See paper for excluded regions in various scenarios. See Figs. 2–6 and Tabs. 14–21 for limits on  $\sigma(ZH^0) \cdot B(H^0 \rightarrow b\bar{b}, \tau^+\tau^-)$  and  $\sigma(H_1^0 H_2^0) \cdot B(H_1^0 H_2^0 \rightarrow b\bar{b}, \tau^+\tau^-)$ .

<sup>3</sup> Search for  $e^+e^- \rightarrow H_1^0 A^0$  in the final states  $b\bar{b}b\bar{b}$  and  $b\bar{b}\tau^+\tau^-$ , and  $e^+e^- \rightarrow H_1^0 Z$ . Universal scalar mass of 1 TeV, SU(2) gaugino mass of 200 GeV, and  $\mu = -200$  GeV are assumed, and two-loop radiative corrections incorporated. The limits hold for  $m_t = 175$  GeV, and for the  $m_h^{\max}$  scenario.

<sup>4</sup> ABBIENDI 04M exclude  $0.7 < \tan\beta < 1.9$ , assuming  $m_t = 174.3$  GeV. Limits for other MSSM benchmark scenarios, as well as for CP violating cases, are also given.

<sup>5</sup> ACHARD 02H also search for the final state  $H_1^0 Z \rightarrow 2A^0 q\bar{q}, A^0 \rightarrow q\bar{q}$ . In addition, the MSSM parameter set in the “large- $\mu$ ” and “no-mixing” scenarios are examined.

<sup>6</sup> HEISTER 02 excludes the range  $0.7 < \tan\beta < 2.3$ . A wider range is excluded with different stop mixing assumptions. Updates BARATE 01c.

<sup>7</sup> KHACHATRYAN 16A search for production of a Higgs boson in gluon fusion and in association with a  $b\bar{b}$  pair followed by the decay  $H_{1,2}^0/A^0 \rightarrow \mu^+\mu^-$  in 5.1 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. See their Fig. 7 for the excluded region in the MSSM parameter space in the  $m_h^{\text{mod+}}$  benchmark scenario and Fig. 9 for limits on cross section times branching ratio.

<sup>8</sup> KHACHATRYAN 15AY search for production of a Higgs boson in association with a  $b$  quark in the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$  in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV and combine with CHATRCHYAN 13AG 7 TeV data. See their Fig. 6 for the limits on cross section times branching ratio for  $m_{A^0} = 100\text{--}900$  GeV and Figs. 7–9 for the excluded region in the MSSM parameter space in various benchmark scenarios.

<sup>9</sup> AAD 14AW search for production of a Higgs boson followed by the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  in 19.5–20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 11 for the limits on cross section times branching ratio and their Figs. 9 and 10 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 140$  GeV, the region  $\tan\beta > 5.4$  is excluded at 95% CL in the  $m_h^{\max}$  scenario.

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# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

- <sup>10</sup> KHACHATRYAN 14M search for production of a Higgs boson in gluon fusion and in association with a  $b$  quark followed by the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  in 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV and 19.7 fb<sup>-1</sup> at  $E_{cm} = 8$  TeV. See their Figs. 7 and 8 for one- and two-dimensional limits on cross section times branching ratio and their Figs. 5 and 6 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 140$  GeV, the region  $\tan\beta > 3.8$  is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>11</sup> AAD 13o search for production of a Higgs boson in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  and  $\mu^+\mu^-$  with 4.7–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. See their Fig. 6 for the excluded region in the MSSM parameter space and their Fig. 7 for the limits on cross section times branching ratio. For  $m_{A^0} = 110$ –170 GeV,  $\tan\beta \gtrsim 10$  is excluded, and for  $\tan\beta = 50$ ,  $m_{A^0}$  below 470 GeV is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>12</sup> AAIJ 13T search for production of a Higgs boson in the forward region in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  in 1.0 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. See their Fig. 2 for the limits on cross section times branching ratio and the excluded region in the MSSM parameter space.
- <sup>13</sup> CHATRCHYAN 13AG search for production of a Higgs boson in association with a  $b$  quark in the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$  in 2.7–4.8 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. See their Fig. 6 for the excluded region in the MSSM parameter space and Fig. 5 for the limits on cross section times branching ratio. For  $m_{A^0} = 90$ –350 GeV, upper bounds on  $\tan\beta$  of 18–42 at 95% CL are obtained in the  $m_h^{\max}$  scenario with  $\mu = +200$  GeV.
- <sup>14</sup> AALTONEN 12AQ combine AALTONEN 12x and ABAZOV 11K. See their Table I and Fig. 1 for the limit on cross section times branching ratio and Fig. 2 for the excluded region in the MSSM parameter space.
- <sup>15</sup> AALTONEN 12x search for associated production of a Higgs boson and a  $b$  quark in the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$ , with 2.6 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. See their Table III and Fig. 15 for the limit on cross section times branching ratio and Figs. 17, 18 for the excluded region in the MSSM parameter space.
- <sup>16</sup> ABAZOV 12c search for production of a Higgs boson in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  with 7.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV and combine with ABAZOV 11w and ABAZOV 11K. See their Figs. 4, 5, and 6 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 90$ –180 GeV,  $\tan\beta \gtrsim 30$  is excluded at 95% CL in the  $m_h^{\max}$  scenario.
- <sup>17</sup> CHATRCHYAN 12K search for production of a Higgs boson in the decay  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$  with 4.6 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. See their Fig. 3 and Table 4 for the excluded region in the MSSM parameter space. For  $m_{A^0} = 160$  GeV, the region  $\tan\beta > 7.1$  is excluded at 95% CL in the  $m_h^{\max}$  scenario. Superseded by KHACHATRYAN 14M.
- <sup>18</sup> ABAZOV 11K search for associated production of a Higgs boson and a  $b$  quark, followed by the decay  $H_{1,2}^0/A^0 \rightarrow b\bar{b}$ , in 5.2 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. See their Fig. 5/ Table 2 for the limit on cross section times branching ratio and Fig. 6 for the excluded region in the MSSM parameter space for  $\mu = -200$  GeV.
- <sup>19</sup> ABAZOV 11w search for associated production of a Higgs boson and a  $b$  quark, followed by the decay  $H_{1,2}^0/A^0 \rightarrow \tau\tau$ , in 7.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. See their Fig. 2 for the limit on cross section times branching ratio and for the excluded region in the MSSM parameter space.
- <sup>20</sup> AALTONEN 09AR search for Higgs bosons decaying to  $\tau^+\tau^-$  in two doublet models in 1.8 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. See their Fig. 2 for the limit on  $\sigma \cdot B(H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-)$  for different Higgs masses, and see their Fig. 3 for the excluded region in the MSSM parameter space.
- <sup>21</sup> ACOSTA 05g search for  $H_{1,2}^0/A^0$  production in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV with  $H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$ . At  $m_{A^0} = 100$  GeV, the obtained cross section upper limit is above theoretical expectation.
- <sup>22</sup> ABBIENDI 03c search for  $e^+e^- \rightarrow H_1^0 Z$  followed by  $H_1^0 \rightarrow A^0 A^0$ ,  $A^0 \rightarrow c\bar{c}$ ,  $g g$ , or  $\tau^+\tau^-$ . In the no-mixing scenario, the region  $m_{H_1^0} = 45$ –85 GeV and  $m_{A^0} = 2$ –9.5 GeV is excluded at 95% CL.
- <sup>23</sup> AKEROYD 02 examine the possibility of a light  $A^0$  with  $\tan\beta < 1$ . Electroweak measurements are found to be inconsistent with such a scenario.

## MASS LIMITS FOR NEUTRAL HIGGS BOSONS IN EXTENDED HIGGS MODELS

This Section covers models which do not fit into either the Standard Model or its simplest minimal Supersymmetric extension (MSSM), leading to anomalous production rates, or nonstandard final states and branching ratios. In particular, this Section covers limits which may apply to generic two-Higgs-doublet models (2HDM), or to special regions of the MSSM parameter space where decays to invisible particles or to photon pairs are dominant (see the review on “Status of Higgs Boson Physics”). Concerning the mass limits for  $H^0$  and  $A^0$  listed below, see the footnotes or the comment lines for details on the nature of the models to which the limits apply.

The observed signal at about 125 GeV, see section “ $H^0$ ”, can be interpreted as one of the neutral Higgs bosons of an extended Higgs sector.

### Mass Limits in General two-Higgs-doublet Models

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
... We do not use the following data for averages, fits, limits, etc. ...				
1	AAD	15BK ATLS	$H_2^0 \rightarrow H^0 H^0$	
2	AAD	15s ATLS	$A^0 \rightarrow Z H^0$	
3	KHACHATRYAN...15BB	CMS	$H_{2,2}^0, A^0 \rightarrow \gamma\gamma$	

4	KHACHATRYAN...15N	CMS	$A^0 \rightarrow Z H^0$	
5	AAD	14M ATLS	$H_2^0 \rightarrow H^\pm W^\mp \rightarrow H^\pm W^\mp W^\mp, H^0 \rightarrow b\bar{b}$	
6	KHACHATRYAN...14Q	CMS	$H_2^0 \rightarrow H^0 H^0, A^0 \rightarrow Z H^0$	
7	AALTONEN	09AR CDF	$p\bar{p} \rightarrow H_{1,2}^0/A^0 + X, H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-$	
8	ABBIENDI	05A OPAL	$H_1^0$ , Type II model	
9	ABDALLAH	05D DLPH	$H^0 \rightarrow 2 \text{ jets}$	
10	ABDALLAH	04o DLPH	$Z \rightarrow f\bar{f} H$	
11	ABDALLAH	04o DLPH	$e^+e^- \rightarrow H^0 Z, H^0 A^0$	
12	ABBIENDI	02D OPAL	$e^+e^- \rightarrow b\bar{b} H$	
13	ABBIENDI	01E OPAL	$H_1^0$ , Type-II model	
14	ABBIENDI	99E OPAL	$\tan\beta > 1$	
15	ABREU	95H DLPH	$Z \rightarrow H^0 Z^*, H^0 A^0$	
16	PICH	92 RVUE	Very light Higgs	
1	AAD 15BK	search for production of a heavy $H_2^0$ decaying to $H^0 H^0$ in the final state $b\bar{b}b\bar{b}$ in 19.5 fb <sup>-1</sup> of $pp$ collisions at $E_{cm} = 8$ TeV. See their Figs. 15–18 for excluded regions in the parameter space.		
2	AAD 15s	search for production of $A^0$ decaying to $Z H^0 \rightarrow \ell^+ \ell^- b\bar{b}, \nu\bar{\nu}b\bar{b}$ and $\ell^+ \ell^- \tau^+ \tau^-$ in 20.3 fb <sup>-1</sup> of $pp$ collisions at $E_{cm} = 8$ TeV. See their Figs. 4 and 5 for excluded regions in the parameter space.		
3	KHACHATRYAN 15BB	search for $H_2^0, A^0 \rightarrow \gamma\gamma$ in 19.7 fb <sup>-1</sup> of $pp$ collisions at $E_{cm} = 8$ TeV. See their Fig. 10 for excluded regions in the two-Higgs-doublet model parameter space.		
4	KHACHATRYAN 15N	search for production of $A^0$ decaying to $Z H^0 \rightarrow \ell^+ \ell^- b\bar{b}$ in 19.7 fb <sup>-1</sup> of $pp$ collisions at $E_{cm} = 8$ TeV. See their Fig. 5 for excluded regions in the $\tan\beta - \cos(\beta - \alpha)$ plane for $m_{A^0} = 300$ GeV.		
5	AAD 14M	search for the decay cascade $H_2^0 \rightarrow H^\pm W^\mp \rightarrow H^0 W^\pm W^\mp, H^0$ decaying to $b\bar{b}$ in 20.3 fb <sup>-1</sup> of $pp$ collisions at $E_{cm} = 8$ TeV. See their Table IV for limits in a two-Higgs-doublet model for $m_{H_2^0} = 325$ –1025 GeV and $m_{H^\pm} = 225$ –825 GeV.		
6	KHACHATRYAN 14Q	search for $H_2^0 \rightarrow H^0 H^0$ and $A^0 \rightarrow Z H^0$ in 19.5 fb <sup>-1</sup> of $pp$ collisions at $E_{cm} = 8$ TeV. See their Figs. 4 and 5 for limits on cross section times branching ratio for $m_{H_2^0} = 260$ –360 GeV and their Figs. 7–9 for limits in two-Higgs-doublet models.		
7	AALTONEN 09AR	search for Higgs bosons decaying to $\tau^+\tau^-$ in two doublet models in 1.8 fb <sup>-1</sup> of $p\bar{p}$ collisions at $E_{cm} = 1.96$ TeV. See their Fig. 2 for the limit on $\sigma \cdot B(H_{1,2}^0/A^0 \rightarrow \tau^+\tau^-)$ for different Higgs masses, and see their Fig. 3 for the excluded region in the MSSM parameter space.		
8	ABBIENDI 05A	search for $e^+e^- \rightarrow H_1^0 A^0$ in general Type-II two-doublet models, with decays $H_1^0, A^0 \rightarrow q\bar{q}, g g, \tau^+\tau^-$ , and $H_1^0 \rightarrow A^0 A^0$ .		
9	ABDALLAH 05D	search for $e^+e^- \rightarrow H^0 Z$ and $H^0 A^0$ with $H^0, A^0$ decaying to two jets of any flavor including $g g$ . The limit is for SM $H^0 Z$ production cross section with $B(H^0 \rightarrow jj) = 1$ .		
10	ABDALLAH 04o	search for $Z \rightarrow b\bar{b} H^0, b\bar{b} A^0, \tau^+\tau^- H^0$ and $\tau^+\tau^- A^0$ in the final states $4b, b\bar{b}\tau^+\tau^-$ , and $4\tau$ . See paper for limits on Yukawa couplings.		
11	ABDALLAH 04o	search for $e^+e^- \rightarrow H^0 Z$ and $H^0 A^0$ , with $H^0, A^0$ decaying to $b\bar{b}, \tau^+\tau^-$ , or $H^0 \rightarrow A^0 A^0$ at $E_{cm} = 189$ –208 GeV. See paper for limits on couplings.		
12	ABBIENDI 02D	search for $Z \rightarrow b\bar{b} H^0$ and $b\bar{b} A^0$ with $H_1^0/A^0 \rightarrow \tau^+\tau^-$ , in the range $4 < m_H < 12$ GeV. See their Fig. 8 for limits on the Yukawa coupling.		
13	ABBIENDI 01E	search for neutral Higgs bosons in general Type-II two-doublet models, at $E_{cm} \leq 189$ GeV. In addition to usual final states, the decays $H_1^0, A^0 \rightarrow q\bar{q}, g g$ are searched for. See their Figs. 15, 16 for excluded regions.		
14	ABBIENDI 99E	search for $e^+e^- \rightarrow H^0 A^0$ and $H^0 Z$ at $E_{cm} = 183$ GeV. The limit is with $m_H = m_A$ in general two Higgs-doublet models. See their Fig. 18 for the exclusion limit in the $m_H$ – $m_A$ plane. Updates the results of ACKERSTAFF 98s.		
15	See Fig. 4 of ABREU 95H	for the excluded region in the $m_{H^0} - m_{A^0}$ plane for general two-doublet models. For $\tan\beta > 1$ , the region $m_{H^0} + m_{A^0} \lesssim 87$ GeV, $m_{H^0} < 47$ GeV is excluded at 95% CL.		
16	PICH 92	analyse $H^0$ with $m_{H^0} < 2m_\mu$ in general two-doublet models. Excluded regions in the space of mass-mixing angles from LEP, beam dump, and $\pi^\pm, \eta$ rare decays are shown in Figs. 3, 4. The considered mass region is not totally excluded.		

### Mass Limits for $H^0$ with Vanishing Yukawa Couplings

These limits assume that  $H^0$  couples to gauge bosons with the same strength as the Standard Model Higgs boson, but has no coupling to quarks and leptons (this is often referred to as “fermiophobic”).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
... We do not use the following data for averages, fits, limits, etc. ...				
1	AALTONEN	13K CDF	$H^0 \rightarrow W W^*$	
2	AALTONEN	13L CDF	$H^0 \rightarrow \gamma\gamma, W W^*, Z Z^*$	
3	AALTONEN	13M TEVA	$H^0 \rightarrow \gamma\gamma, W W^*, Z Z^*$	
4	ABAZOV	13G D0	$H^0 \rightarrow W W^*$	
5	ABAZOV	13H D0	$H^0 \rightarrow \gamma\gamma$	
6	ABAZOV	13I D0	$H^0 \rightarrow W W^*$	
7	ABAZOV	13J D0	$H^0 \rightarrow W W^*, Z Z^*$	
8	ABAZOV	13L D0	$H^0 \rightarrow \gamma\gamma, W W^*, Z Z^*$	
9	CHATRCHYAN 13AL	CMS	$H^0 \rightarrow \gamma\gamma$	
10	AAD	12N ATLS	$H^0 \rightarrow \gamma\gamma$	
11	AALTONEN	12AN CDF	$H^0 \rightarrow \gamma\gamma$	

# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

none 110–194	95	12	CHATRCHYAN12A0	CMS	$H^0 \rightarrow \gamma\gamma, WW^{(*)}, ZZ^{(*)}$
none 70–106	95	13	AALTONEN	09AB CDF	$H^0 \rightarrow \gamma\gamma$
none 70–100	95	14	ABAZOV	08U D0	$H^0 \rightarrow \gamma\gamma$
>105.8	95	15	SCHAEEL	07 ALEP	$e^+e^- \rightarrow H^0 Z, H^0 \rightarrow WW^{*}$
>104.1	95	16,17	ABDALLAH	04L DLPH	$e^+e^- \rightarrow H^0 Z, H^0 \rightarrow \gamma\gamma$
>107	95	18	ACHARD	03C L3	$H^0 \rightarrow WW^{*}, ZZ^{*}, \gamma\gamma$
>105.5	95	16,19	ABBIENDI	02F OPAL	$H^0 \rightarrow \gamma\gamma$
>105.4	95	20	ACHARD	02C L3	$H^0 \rightarrow \gamma\gamma$
none 60–82	95	21	AFFOLDER	01H CDF	$p\bar{p} \rightarrow H^0 W/Z, H^0 \rightarrow \gamma\gamma$
> 94.9	95	22	ACCIARRI	00S L3	$e^+e^- \rightarrow H^0 Z, H^0 \rightarrow \gamma\gamma$
>100.7	95	23	BARATE	00L ALEP	$e^+e^- \rightarrow H^0 Z, H^0 \rightarrow \gamma\gamma$
> 96.2	95	24	ABBIENDI	99O OPAL	$e^+e^- \rightarrow H^0 Z, H^0 \rightarrow \gamma\gamma$
> 78.5	95	25	ABBOTT	99B D0	$p\bar{p} \rightarrow H^0 W/Z, H^0 \rightarrow \gamma\gamma$
		26	ABREU	99P DLPH	$e^+e^- \rightarrow H^0 \gamma$ and/or $H^0 \rightarrow \gamma\gamma$

- 1 AALTONEN 13K search for  $H^0 \rightarrow WW^{(*)}$  in 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which corresponds to (1.3–6.6) times the expected cross section is given in the range  $m_{H^0} = 110$ –200 GeV at 95% CL.
- 2 AALTONEN 13L combine all CDF searches with 9.45–10.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV.
- 3 AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV.
- 4 ABAZOV 13G search for  $H^0 \rightarrow WW^{(*)}$  in 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which corresponds to (2–9) times the expected cross section is given for  $m_{H^0} = 100$ –200 GeV at 95% CL.
- 5 ABAZOV 13H search for  $H^0 \rightarrow \gamma\gamma$  in 9.6 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV.
- 6 ABAZOV 13I search for  $H^0$  production in the final state with one lepton and two or more jets plus missing  $E_T$  in 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The search is sensitive to  $WH^0, ZH^0$  and vector-boson fusion Higgs production with  $H^0 \rightarrow WW^{(*)}$ . A limit on cross section times branching ratio which corresponds to (8–30) times the expected cross section is given in the range  $m_{H^0} = 100$ –200 GeV at 95% CL.
- 7 ABAZOV 13J search for  $H^0$  production in the final states  $e\mu\mu, e\mu\mu, \mu\tau\tau$ , and  $e^\pm\mu^\pm$  in 8.6–9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The search is sensitive to  $WH^0, ZH^0$  production with  $H^0 \rightarrow WW^{(*)}, ZZ^{(*)}$ , decaying to leptonic final states. A limit on cross section times branching ratio which corresponds to (2.4–13.0) times the expected cross section is given in the range  $m_{H^0} = 100$ –200 GeV at 95% CL.
- 8 ABAZOV 13L combine all D0 results with up to 9.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV.
- 9 CHATRCHYAN 13AL search for  $H^0 \rightarrow \gamma\gamma$  in 5.1 fb<sup>-1</sup> and 5.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  and 8 TeV.
- 10 AAD 12N search for  $H^0 \rightarrow \gamma\gamma$  with 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 110$ –150 GeV.
- 11 AALTONEN 12AN search for  $H^0 \rightarrow \gamma\gamma$  with 10 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV in the mass range  $m_{H^0} = 100$ –150 GeV.
- 12 CHATRCHYAN 12A0 use data from CHATRCHYAN 12G, CHATRCHYAN 12E, CHATRCHYAN 12H, CHATRCHYAN 12I, CHATRCHYAN 12D, and CHATRCHYAN 12C.
- 13 AALTONEN 09AB search for  $H^0 \rightarrow \gamma\gamma$  in 3.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV in the mass range  $m_{H^0} = 70$ –150 GeV. Associated  $H^0 W, H^0 Z$  production and  $WW, ZZ$  fusion are considered.
- 14 ABAZOV 08U search for  $H^0 \rightarrow \gamma\gamma$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV in the mass range  $m_{H^0} = 70$ –150 GeV. Associated  $H^0 W, H^0 Z$  production and  $WW, ZZ$  fusion are considered. See their Tab. 1 for the limit on  $\sigma \cdot B(H^0 \rightarrow \gamma\gamma)$ , and see their Fig. 3 for the excluded region in the  $m_{H^0} - B(H^0 \rightarrow \gamma\gamma)$  plane.
- 15 SCHAEEL 07 search for Higgs bosons in association with a fermion pair and decaying to  $WW^{*}$ . The limit is from this search and HEISTER 02L for a  $H^0$  with SM production cross section.
- 16 Search for associated production of a  $\gamma\gamma$  resonance with a  $Z$  boson, followed by  $Z \rightarrow q\bar{q}, \ell^+\ell^-$ , or  $\nu\bar{\nu}$ , at  $E_{\text{cm}} \leq 209$  GeV. The limit is for a  $H^0$  with SM production cross section.
- 17 Updates ABREU 01F.
- 18 ACHARD 03C search for  $e^+e^- \rightarrow ZH^0$  followed by  $H^0 \rightarrow WW^{*}$  or  $ZZ^{*}$  at  $E_{\text{cm}} = 200$ –209 GeV and combine with the ACHARD 02C result. The limit is for a  $H^0$  with SM production cross section. For  $B(H^0 \rightarrow WW^{*}) + B(H^0 \rightarrow ZZ^{*}) = 1$ ,  $m_{H^0} > 108.1$  GeV is obtained. See fig. 6 for the limits under different BR assumptions.
- 19 For  $B(H^0 \rightarrow \gamma\gamma) = 1$ ,  $m_{H^0} > 117$  GeV is obtained.
- 20 ACHARD 02C search for associated production of a  $\gamma\gamma$  resonance with a  $Z$  boson, followed by  $Z \rightarrow q\bar{q}, \ell^+\ell^-$ , or  $\nu\bar{\nu}$ , at  $E_{\text{cm}} \leq 209$  GeV. The limit is for a  $H^0$  with SM production cross section. For  $B(H^0 \rightarrow \gamma\gamma) = 1$ ,  $m_{H^0} > 114$  GeV is obtained.
- 21 AFFOLDER 01H search for associated production of a  $\gamma\gamma$  resonance and a  $W$  or  $Z$  (tagged by two jets, an isolated lepton, or missing  $E_T$ ). The limit assumes Standard Model values for the production cross section and for the couplings of the  $H^0$  to  $W$  and  $Z$  bosons. See their Fig. 11 for limits with  $B(H^0 \rightarrow \gamma\gamma) < 1$ .
- 22 ACCIARRI 00S search for associated production of a  $\gamma\gamma$  resonance with a  $q\bar{q}, \nu\bar{\nu}$ , or  $\ell^+\ell^-$  pair in  $e^+e^-$  collisions at  $E_{\text{cm}} = 189$  GeV. The limit is for a  $H^0$  with SM production cross section. For  $B(H^0 \rightarrow \gamma\gamma) = 1$ ,  $m_{H^0} > 98$  GeV is obtained. See their Fig. 5 for limits on  $B(H \rightarrow \gamma\gamma) \cdot \sigma(e^+e^- \rightarrow Hf\bar{f}) / \sigma(e^+e^- \rightarrow Hf\bar{f})$  (SM).
- 23 BARATE 00L search for associated production of a  $\gamma\gamma$  resonance with a  $q\bar{q}, \nu\bar{\nu}$ , or  $\ell^+\ell^-$  pair in  $e^+e^-$  collisions at  $E_{\text{cm}} = 88$ –202 GeV. The limit is for a  $H^0$  with SM production cross section. For  $B(H^0 \rightarrow \gamma\gamma) = 1$ ,  $m_{H^0} > 109$  GeV is obtained. See their Fig. 3 for limits on  $B(H \rightarrow \gamma\gamma) \cdot \sigma(e^+e^- \rightarrow Hf\bar{f}) / \sigma(e^+e^- \rightarrow Hf\bar{f})$  (SM).
- 24 ABBIENDI 99O search for associated production of a  $\gamma\gamma$  resonance with a  $q\bar{q}, \nu\bar{\nu}$ , or  $\ell^+\ell^-$  pair in  $e^+e^-$  collisions at 189 GeV. The limit is for a  $H^0$  with SM production cross

section. See their Fig. 4 for limits on  $\sigma(e^+e^- \rightarrow H^0 Z^0) \times B(H^0 \rightarrow \gamma\gamma) \times B(X^0 \rightarrow f\bar{f})$  for various masses. Updates the results of ACKERSTAFF 98Y.

- 25 ABBOTT 99B search for associated production of a  $\gamma\gamma$  resonance and a dijet pair. The limit assumes Standard Model values for the production cross section and for the couplings of the  $H^0$  to  $W$  and  $Z$  bosons. Limits in the range of  $\sigma(H^0 + Z/W) \cdot B(H^0 \rightarrow \gamma\gamma) = 0.80$ –0.34 pb are obtained in the mass range  $m_{H^0} = 65$ –150 GeV.

- 26 ABREU 99P search for  $e^+e^- \rightarrow H^0 \gamma$  with  $H^0 \rightarrow b\bar{b}$  or  $\gamma\gamma$ , and  $e^+e^- \rightarrow H^0 q\bar{q}$  with  $H^0 \rightarrow \gamma\gamma$ . See their Fig. 4 for limits on  $\sigma \times B$ . Explicit limits within an effective interaction framework are also given.

## Mass Limits for $H^0$ Decaying to Invisible Final States

These limits are for a neutral scalar  $H^0$  which predominantly decays to invisible final states. Standard Model values are assumed for the couplings of  $H^0$  to ordinary particles unless otherwise stated.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
		1 AAD	15BD ATLS	$p\bar{p} \rightarrow H^0 WX, H^0 ZX$
		2 AAD	15BH ATLS	jet + missing $E_T$
		3 AAD	14BA ATLS	secondary vertex
		4 AAD	14O ATLS	$p\bar{p} \rightarrow H^0 ZX$
		5 CHATRCHYAN14B	CMS	$p\bar{p} \rightarrow H^0 ZX, qqH^0 X$
		6 AAD	13AG ATLS	secondary vertex
		7 AAD	13AT ATLS	electron jets
		8 CHATRCHYAN13BJ	CMS	
		9 AAD	12AQ ATLS	secondary vertex
		10 AALTONEN	12AB CDF	secondary vertex
		11 AALTONEN	12U CDF	secondary vertex
>108.2	95	12 ABBIENDI	10 OPAL	
		13 ABBIENDI	07 OPAL	large width
>112.3	95	14 ACHARD	05 L3	
>112.1	95	14 ABDALLAH	04B DLPH	
>114.1	95	14 HEISTER	02 ALEP	$E_{\text{cm}} \leq 209$ GeV
>106.4	95	14 BARATE	01C ALEP	$E_{\text{cm}} \leq 202$ GeV
> 89.2	95	15 ACCIARRI	00M L3	

- 1 AAD 15BD search for  $p\bar{p} \rightarrow H^0 WX$  and  $p\bar{p} \rightarrow H^0 ZX$  with  $W$  or  $Z$  decaying hadronically and  $H^0$  decaying to invisible final states in 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. See their Fig. 6 for a limit on the cross section times branching ratio for  $m_{H^0} = 115$ –300 GeV.
- 2 AAD 15BH search for events with a jet and missing  $E_T$  in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. Limits on  $\sigma(H^0) B(H^0 \rightarrow \text{invisible}) < (44\text{--}10)$  pb (95%CL) is given for  $m_{H^0} = 115$ –300 GeV.
- 3 AAD 14BA search for  $H^0$  production in the decay mode  $H^0 \rightarrow X^0 X^0$ , where  $X^0$  is a long-lived particle which decays to collimated pairs of  $e^+e^-, \mu^+\mu^-,$  or  $\pi^+\pi^-$  plus invisible particles, in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Figs. 15 and 16 for limits on cross section times branching ratio.
- 4 AAD 14O search for  $p\bar{p} \rightarrow H^0 ZX, Z \rightarrow \ell\ell$  with  $H^0$  decaying to invisible final states in 4.5 fb<sup>-1</sup> at  $E_{\text{cm}} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $E_{\text{cm}} = 8$  TeV. See their Fig. 3 for a limit on the cross section times branching ratio for  $m_{H^0} = 110$ –400 GeV.
- 5 CHATRCHYAN 14B search for  $p\bar{p} \rightarrow H^0 ZX, Z \rightarrow \ell\ell$  and  $Z \rightarrow b\bar{b}$ , and also  $p\bar{p} \rightarrow qqH^0 X$  with  $H^0$  decaying to invisible final states using data at  $E_{\text{cm}} = 7$  and 8 TeV. See their Figs. 10, 11 for limits on the cross section times branching ratio for  $m_{H^0} = 100$ –400 GeV.
- 6 AAD 13AG search for  $H^0$  production in the decay mode  $H^0 \rightarrow X^0 X^0$ , where  $X^0$  is a long-lived particle which decays to  $\mu^+\mu^- X^0$ , in 1.9 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 7 for limits on cross section times branching ratio.
- 7 AAD 13AT search for  $H^0$  production in the decay  $H^0 \rightarrow X^0 X^0$ , where  $X^0$  eventually decays to clusters of collimated  $e^+e^-$  pairs, in 2.04 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 3 for limits on cross section times branching ratio.
- 8 CHATRCHYAN 13BJ search for  $H^0$  production in the decay chain  $H^0 \rightarrow X^0 X^0, X^0 \rightarrow \mu^+\mu^- X^0$  in 5.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 2 for limits on cross section times branching ratio.
- 9 AAD 12AQ search for  $H^0$  production in the decay mode  $H^0 \rightarrow X^0 X^0$ , where  $X^0$  is a long-lived particle which decays mainly to  $b\bar{b}$  in the muon detector, in 1.94 fb<sup>-1</sup> of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 3 for limits on cross section times branching ratio for  $m_{H^0} = 120, 140$  GeV,  $m_{X^0} = 20, 40$  GeV in the  $c\tau$  range of 0.5–35 m.
- 10 AALTONEN 12AB search for  $H^0$  production in the decay  $H^0 \rightarrow X^0 X^0$ , where  $X^0$  eventually decays to clusters of collimated  $\ell^+\ell^-$  pairs, in 5.1 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. Cross section limits are provided for a benchmark MSSM model incorporating the parameters given in Table VI.
- 11 AALTONEN 12U search for  $H^0$  production in the decay mode  $H^0 \rightarrow X^0 X^0$ , where  $X^0$  is a long-lived particle with  $c\tau \approx 1$  cm which decays mainly to  $b\bar{b}$ , in 3.2 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. See their Figs. 9 and 10 for limits on cross section times branching ratio for  $m_{H^0} = (130\text{--}170)$  GeV,  $m_{X^0} = 20, 40$  GeV.
- 12 ABBIENDI 10 search for  $e^+e^- \rightarrow H^0 Z$  with  $H^0$  decaying invisibly. The limit assumes SM production cross section and  $B(H^0 \rightarrow \text{invisible}) = 1$ .
- 13 ABBIENDI 07 search for  $e^+e^- \rightarrow H^0 Z$  with  $Z \rightarrow q\bar{q}$  and  $H^0$  decaying to invisible final states. The  $H^0$  width is varied between 1 GeV and 3 TeV. A limit  $\sigma \cdot B(H^0 \rightarrow \text{invisible}) < (0.07\text{--}0.57)$  pb (95%CL) is obtained at  $E_{\text{cm}} = 206$  GeV for  $m_{H^0} = 60$ –114 GeV.
- 14 Search for  $e^+e^- \rightarrow H^0 Z$  with  $H^0$  decaying invisibly. The limit assumes SM production cross section and  $B(H^0 \rightarrow \text{invisible}) = 1$ .
- 15 ACCIARRI 00M search for  $e^+e^- \rightarrow ZH^0$  with  $H^0$  decaying invisibly at  $E_{\text{cm}} = 183$ –189 GeV. The limit assumes SM production cross section and  $B(H^0 \rightarrow \text{invisible}) = 1$ . See their Fig. 6 for limits for smaller branching ratios.



See key on page 601

# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

### Mass Limits for Light $A^0$

These limits are for a pseudoscalar  $A^0$  in the mass range below  $\mathcal{O}(10)$  GeV.

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1	KHACHATRYAN 16F	CMS	$H^0 \rightarrow A^0 A^0$
2	LEES 15H	BABR	$\Upsilon(1S) \rightarrow A^0 \gamma$
3	LEES 13C	BABR	$\Upsilon(1S) \rightarrow A^0 \gamma$
4	LEES 13L	BABR	$\Upsilon(1S) \rightarrow A^0 \gamma$
5	LEES 13R	BABR	$\Upsilon(1S) \rightarrow A^0 \gamma$
6	CHATRCHYAN 12V	CMS	$A^0 \rightarrow \mu^+ \mu^-$
7	AALTONEN 11P	CDF	$t \rightarrow b H^+, H^+ \rightarrow W^+ A^0$
8,9	ABOUZAID 11A	KTEV	$K_L \rightarrow \pi^0 \pi^0 A^0, A^0 \rightarrow \mu^+ \mu^-$
10	DEL-AMO-SANCHEZ 11J	BABR	$\Upsilon(1S) \rightarrow A^0 \gamma$
11	LEES 11H	BABR	$\Upsilon(2S, 3S) \rightarrow A^0 \gamma$
12	ANDREAS 10	RVUE	
9,13	HYUN 10	BELL	$B^0 \rightarrow K^{*0} A^0, A^0 \rightarrow \mu^+ \mu^-$
9,14	HYUN 10	BELL	$B^0 \rightarrow \rho^0 A^0, A^0 \rightarrow \mu^+ \mu^-$
15	AUBERT 09P	BABR	$\Upsilon(3S) \rightarrow A^0 \gamma$
16	AUBERT 09Z	BABR	$\Upsilon(2S) \rightarrow A^0 \gamma$
17	AUBERT 09Z	BABR	$\Upsilon(3S) \rightarrow A^0 \gamma$
9,18	TUNG 09	K391	$K_L \rightarrow \pi^0 \pi^0 A^0, A^0 \rightarrow \gamma \gamma$
19	LOVE 08	CLEO	$\Upsilon(1S) \rightarrow A^0 \gamma$
20	BESSION 07	CLEO	$\Upsilon(1S) \rightarrow \eta_b \gamma$
21	PARK 05	HYCP	$\Sigma^+ \rightarrow p A^0, A^0 \rightarrow \mu^+ \mu^-$
22	BALEST 95	CLE2	$\Upsilon(1S) \rightarrow A^0 \gamma$
23	ANTREASIAN 90c	CBAL	$\Upsilon(1S) \rightarrow A^0 \gamma$

- 1 KHACHATRYAN 16F search for the decay  $H^0 \rightarrow A^0 A^0 \rightarrow \tau^+ \tau^- \tau^+ \tau^-$  in 19.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 8 for cross section limits for  $m_{A^0} = 4-8$  GeV.
- 2 LEES 15H search for the process  $\Upsilon(2S) \rightarrow \Upsilon(1S) \pi^+ \pi^- \rightarrow A^0 \gamma \pi^+ \pi^-$  with  $A^0$  decaying to  $c\bar{c}$  and give limits on  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow c\bar{c})$  in the range  $7.4 \times 10^{-5} - 2.4 \times 10^{-3}$  (90% CL) for  $4.00 \leq m_{A^0} \leq 8.95$  and  $9.10 \leq m_{A^0} \leq 9.25$  GeV. See their Fig. 6.
- 3 LEES 13C search for the process  $\Upsilon(2S, 3S) \rightarrow \Upsilon(1S) \pi^+ \pi^- \rightarrow A^0 \gamma \pi^+ \pi^-$  with  $A^0$  decaying to  $\mu^+ \mu^-$  and give limits on  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \mu^+ \mu^-)$  in the range  $(0.3-9.7) \times 10^{-6}$  (90% CL) for  $0.212 \leq m_{A^0} \leq 9.20$  GeV. See their Fig. 5(e) for limits on the  $b-A^0$  Yukawa coupling derived by combining this result with AUBERT 09Z.
- 4 LEES 13L search for the process  $\Upsilon(2S) \rightarrow \Upsilon(1S) \pi^+ \pi^- \rightarrow A^0 \gamma \pi^+ \pi^-$  with  $A^0$  decaying to  $g\bar{g}$  or  $s\bar{s}$  and give limits on  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow g\bar{g})$  between  $1 \times 10^{-6}$  and  $2 \times 10^{-2}$  (90% CL) for  $0.5 \leq m_{A^0} \leq 9.0$  GeV, and  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow s\bar{s})$  between  $4 \times 10^{-6}$  and  $1 \times 10^{-3}$  (90% CL) for  $1.5 \leq m_{A^0} \leq 9.0$  GeV. See their Fig. 4.
- 5 LEES 13R search for the process  $\Upsilon(2S) \rightarrow \Upsilon(1S) \pi^+ \pi^- \rightarrow A^0 \gamma \pi^+ \pi^-$  with  $A^0$  decaying to  $\tau^+ \tau^-$  and give limits on  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \tau^+ \tau^-)$  in the range  $0.9-13 \times 10^{-5}$  (90% CL) for  $3.6 \leq m_{A^0} \leq 9.2$  GeV. See their Fig. 4 for limits on the  $b-A^0$  Yukawa coupling derived by combining this result with AUBERT 09P.
- 6 CHATRCHYAN 12V search for  $A^0$  production in the decay  $A^0 \rightarrow \mu^+ \mu^-$  with 1.3 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on  $\sigma(A^0) \cdot B(A^0 \rightarrow \mu^+ \mu^-)$  in the range  $(1.5-7.5)$  pb is given for  $m_{A^0} = (5.5-8.7)$  and  $(11.5-14)$  GeV at 95% CL.
- 7 AALTONEN 11P search in 2.7 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV for the decay chain  $t \rightarrow b H^+, H^+ \rightarrow W^+ A^0, A^0 \rightarrow \tau^+ \tau^-$  with  $m_{A^0}$  between 4 and 9 GeV. See their Fig. 4 for limits on  $B(t \rightarrow b H^+)$  for  $90 < m_{H^+} < 160$  GeV.
- 8 ABOUZAID 11A search for the decay chain  $K_L \rightarrow \pi^0 \pi^0 A^0, A^0 \rightarrow \mu^+ \mu^-$  and give a limit  $B(K_L \rightarrow \pi^0 \pi^0 A^0) \cdot B(A^0 \rightarrow \mu^+ \mu^-) < 1.0 \times 10^{-10}$  at 90% CL for  $m_{A^0} = 214.3$  MeV. The search was motivated by PARK 05.
- 9 DEL-AMO-SANCHEZ 11J search for the process  $\Upsilon(2S) \rightarrow \Upsilon(1S) \pi^+ \pi^- \rightarrow A^0 \gamma \pi^+ \pi^-$  with  $A^0$  decaying to invisible final states. They give limits on  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \text{invisible})$  in the range  $(1.9-4.5) \times 10^{-6}$  (90% CL) for  $0 \leq m_{A^0} \leq 8.0$  GeV, and  $(2.7-37) \times 10^{-6}$  for  $8.0 \leq m_{A^0} \leq 9.2$  GeV.
- 11 LEES 11H search for the process  $\Upsilon(2S, 3S) \rightarrow A^0 \gamma$  with  $A^0$  decaying hadronically and give limits on  $B(\Upsilon(2S, 3S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \text{hadrons})$  in the range  $1 \times 10^{-6} - 8 \times 10^{-5}$  (90% CL) for  $0.3 < m_{A^0} < 7$  GeV. The decay rates for  $\Upsilon(2S)$  and  $\Upsilon(3S)$  are assumed to be equal up to the phase space factor. See their Fig. 5.
- 12 ANDREAS 10 analyze constraints from rare decays and other processes on a light  $A^0$  with  $m_{A^0} < 2m_\mu$  and give limits on its coupling to fermions at the level of  $10^{-4}$  times the Standard Model value.
- 13 HYUN 10 search for the decay chain  $B^0 \rightarrow K^{*0} A^0, A^0 \rightarrow \mu^+ \mu^-$  and give a limit on  $B(B^0 \rightarrow K^{*0} A^0) \cdot B(A^0 \rightarrow \mu^+ \mu^-)$  in the range  $(2.26-5.53) \times 10^{-8}$  at 90% CL for  $m_{A^0} = 212-300$  MeV. The limit for  $m_{A^0} = 214.3$  MeV is  $2.26 \times 10^{-8}$ .
- 14 HYUN 10 search for the decay chain  $B^0 \rightarrow \rho^0 A^0, A^0 \rightarrow \mu^+ \mu^-$  and give a limit on  $B(B^0 \rightarrow \rho^0 A^0) \cdot B(A^0 \rightarrow \mu^+ \mu^-)$  in the range  $(1.73-4.51) \times 10^{-8}$  at 90% CL for  $m_{A^0} = 212-300$  MeV. The limit for  $m_{A^0} = 214.3$  MeV is  $1.73 \times 10^{-8}$ .
- 15 AUBERT 09P search for the process  $\Upsilon(3S) \rightarrow A^0 \gamma$  with  $A^0 \rightarrow \tau^+ \tau^-$  for  $4.03 < m_{A^0} < 9.52$  and  $9.61 < m_{A^0} < 10.10$  GeV, and give limits on  $B(\Upsilon(3S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \tau^+ \tau^-)$  in the range  $(1.5-16) \times 10^{-5}$  (90% CL).
- 16 AUBERT 09Z search for the process  $\Upsilon(2S) \rightarrow A^0 \gamma$  with  $A^0 \rightarrow \mu^+ \mu^-$  for  $0.212 < m_{A^0} < 9.3$  GeV and give limits on  $B(\Upsilon(2S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \mu^+ \mu^-)$  in the range  $(0.3-8) \times 10^{-6}$  (90% CL).
- 17 AUBERT 09Z search for the process  $\Upsilon(3S) \rightarrow A^0 \gamma$  with  $A^0 \rightarrow \mu^+ \mu^-$  for  $0.212 < m_{A^0} < 9.3$  GeV and give limits on  $B(\Upsilon(3S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \mu^+ \mu^-)$  in the range  $(0.3-5) \times 10^{-6}$  (90% CL).

- 18 TUNG 09 search for the decay chain  $K_L \rightarrow \pi^0 \pi^0 A^0, A^0 \rightarrow \gamma \gamma$  and give a limit on  $B(K_L \rightarrow \pi^0 \pi^0 A^0) \cdot B(A^0 \rightarrow \gamma \gamma)$  in the range  $(2.4-10.7) \times 10^{-7}$  at 90% CL for  $m_{A^0} = 194.3-219.3$  MeV. The limit for  $m_{A^0} = 214.3$  MeV is  $2.4 \times 10^{-7}$ .
- 19 LOVE 08 search for the process  $\Upsilon(1S) \rightarrow A^0 \gamma$  with  $A^0 \rightarrow \mu^+ \mu^-$  (for  $m_{A^0} < 2m_\tau$ ) and  $A^0 \rightarrow \tau^+ \tau^-$ . Limits on  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot B(A^0 \rightarrow \ell^+ \ell^-)$  in the range  $10^{-6}-10^{-4}$  (90% CL) are given.
- 20 BESSION 07 give a limit  $B(\Upsilon(1S) \rightarrow \eta_b \gamma) \cdot B(\eta_b \rightarrow \tau^+ \tau^-) < 0.27\%$  (95% CL), which constrains a possible  $A^0$  exchange contribution to the  $\eta_b$  decay.
- 21 PARK 05 found three candidate events for  $\Sigma^+ \rightarrow p \mu^+ \mu^-$  in the HyperCP experiment. Due to a narrow spread in dimuon mass, they hypothesize the events as a possible signal of a new boson. It can be interpreted as a neutral particle with  $m_{A^0} = 214.3 \pm 0.5$  MeV and the branching fraction  $B(\Sigma^+ \rightarrow p A^0) \cdot B(A^0 \rightarrow \mu^+ \mu^-) = (3.1^{+2.4}_{-1.9} \pm 1.5) \times 10^{-8}$ .
- 22 BALEST 95 give limits  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot 1.5 \times 10^{-5}$  at 90% CL for  $m_{A^0} < 5$  GeV. The limit becomes  $< 10^{-4}$  for  $m_{A^0} < 7.7$  GeV.
- 23 ANTREASIAN 90c give limits  $B(\Upsilon(1S) \rightarrow A^0 \gamma) \cdot 5.6 \times 10^{-5}$  at 90% CL for  $m_{A^0} < 7.2$  GeV.  $A^0$  is assumed not to decay in the detector.

### Other Mass Limits

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1		AAD 16C	ATLS	$H^0 \rightarrow W^+ W^-$
2		KHACHATRYAN 16F	CMS	$H^0_2 \rightarrow H^0_1 H^0_1$
3		AAD 15BK	ATLS	$H^0_2 \rightarrow H^0_1 H^0_1$
4		AAD 15BZ	ATLS	$H^0_2 \rightarrow A^0 A^0$
5		AAD 15BZ	ATLS	$H^0_2 \rightarrow A^0 A^0$
6		AAD 15CE	ATLS	$H^0_2 \rightarrow H^0 H^0$
7		AAD 15H	ATLS	$H^0_2 \rightarrow H^0 H^0$
8		AAD 15S	ATLS	$A^0 \rightarrow Z H^0$
9		KHACHATRYAN 15AW	CMS	$H^0_2 \rightarrow W^+ W^-, Z Z$
10		KHACHATRYAN 15BB	CMS	$H^0 \rightarrow \gamma \gamma$
11		KHACHATRYAN 15N	CMS	$A^0 \rightarrow Z H^0$
12		KHACHATRYAN 15O	CMS	$A^0 \rightarrow Z H^0$
13		AAD 14AP	ATLS	$H^0 \rightarrow H^\pm \tau^\mp$
14		AAD 14M	ATLS	$H^0_2 \rightarrow H^\pm W^\mp \rightarrow H^0 W^\pm W^\mp, H^0 \rightarrow b\bar{b}$
15		CHATRCHYAN 14G	CMS	$H^0 \rightarrow W W^{(*)}$
16		KHACHATRYAN 14P	CMS	$H^0 \rightarrow \gamma \gamma$
17		AALTONEN 13P	CDF	$H^0 \rightarrow H^\pm W^\mp \rightarrow H^0 W^\pm W^\mp$
18		CHATRCHYAN 13BJ	CMS	$H^0 \rightarrow A^0 A^0$
19		AALTONEN 11P	CDF	$t \rightarrow b H^+, H^+ \rightarrow W^+ A^0$
20		ABBIENDI 10	OPAL	$H^0 \rightarrow \tilde{\chi}^0_1 \tilde{\chi}^0_2$
21		SCHAELE 10	ALEP	$H^0 \rightarrow A^0 A^0$
22		ABAZOV 09V	D0	$H^0 \rightarrow A^0 A^0$
23		ABBIENDI 05A	OPAL	$A^0$ , Type II model
24		ABBIENDI 04K	OPAL	$H^0 \rightarrow 2$ jets
25		ABDALLAH 04	DLPH	$H^0 V V$ couplings
26		ACHARD 04B	L3	$H^0 \rightarrow 2$ jets
27		ACHARD 04F	L3	Anomalous coupling
28		ABBIENDI 03F	OPAL	$e^+ e^- \rightarrow H^0 Z, H^0 \rightarrow \text{any}$
29		ABBIENDI 03G	OPAL	$H^0 \rightarrow A^0 A^0$
30,31		HEISTER 02L	ALEP	$H^0 \rightarrow \gamma \gamma$
32		HEISTER 02M	ALEP	$H^0 \rightarrow 2$ jets or $\tau^+ \tau^-$
33		ABBIENDI 01E	OPAL	$A^0$ , Type-II model
34		ACCIARRI 00R	L3	$e^+ e^- \rightarrow H^0 \gamma$ and/or $H^0 \rightarrow \gamma \gamma$
35		ACCIARRI 00R	L3	$e^+ e^- \rightarrow e^+ e^- H^0$
36		GONZALEZ-G. 98B	RVUE	Anomalous coupling
37		KRAWCZYK 97	RVUE	$(g-2)_\mu$
38		ALEXANDER 96H	OPAL	$Z \rightarrow H^0 \gamma$

- 1 AAD 16C search for production of a heavy  $H^0$  state decaying to  $W^+ W^-$  in the final states  $\ell\nu\ell\nu$  and  $\ell\nu q\bar{q}$  in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Figs. 12, 13, and 16 for upper limits on  $\sigma(H^0) \cdot B(H^0 \rightarrow W^+ W^-)$  for  $m_{H^0}$  ranging from 300 GeV to 1000 or 1500 GeV with various assumptions on the total width of  $H^0$ .
- 2 KHACHATRYAN 16F search for the decay  $H^0_2 \rightarrow H^0_1 H^0_1 \rightarrow \tau^+ \tau^- \tau^+ \tau^-$  with  $m_{H^0_2} = 125$  GeV in 19.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 8 for cross section limits for  $m_{H^0_1} = 4-8$  GeV.
- 3 AAD 15BK search for production of a heavy  $H^0_2$  decaying to  $H^0 H^0$  in the final state  $b\bar{b}b\bar{b}$  in 19.5 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 14(c) for  $\sigma(H^0_2) \cdot B(H^0_2 \rightarrow H^0 H^0)$  for  $m_{H^0_2} = 500-1500$  GeV with  $\Gamma_{H^0_2} = 1$  GeV.
- 4 AAD 15BZ search for the decay  $H^0 \rightarrow A^0 A^0 \rightarrow \mu^+ \mu^- \tau^+ \tau^-$  ( $m_{H^0} = 125$  GeV) in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 6 for limits on cross section times branching ratio for  $m_{A^0} = 3.7-5.0$  GeV.
- 5 AAD 15BZ search for a state  $H^0_2$  via the decay  $H^0_2 \rightarrow A^0 A^0 \rightarrow \mu^+ \mu^- \tau^+ \tau^-$  in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV. See their Fig. 6 for limits on cross section times branching ratio for  $m_{H^0_2} = 100-500$  GeV and  $m_{A^0} = 5$  GeV.

# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

- <sup>6</sup> AAD 15CE search for production of a heavy  $H_2^0$  decaying to  $H^0 H^0$  in the final states  $b\bar{b}\tau^+\tau^-$  and  $\gamma\gamma WW^*$  in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV and combine with data from AAD 15H and AAD 15BK. A limit  $\sigma(H_2^0) B(H_2^0 \rightarrow H^0 H^0) < 2.1\text{--}0.011$  pb (95% CL) is given for  $m_{H_2^0} = 260\text{--}1000$  GeV. See their Fig. 6.
- <sup>7</sup> AAD 15H search for production of a heavy  $H_2^0$  decaying to  $H^0 H^0$  in the final state  $\gamma\gamma b\bar{b}$  in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. A limit of  $\sigma(H_2^0) B(H_2^0 \rightarrow H^0 H^0) < 3.5\text{--}0.7$  pb is given for  $m_{H_2^0} = 260\text{--}500$  GeV at 95% CL. See their Fig. 3.
- <sup>8</sup> AAD 15s search for production of  $A^0$  decaying to  $ZH^0 \rightarrow \ell^+\ell^-b\bar{b}, \nu\bar{\nu}b\bar{b}$  and  $\ell^+\ell^-\tau^+\tau^-$  in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Fig. 3 for cross section limits for  $m_{A^0} = 200\text{--}1000$  GeV.
- <sup>9</sup> KHACHATRYAN 15AW search for production of a heavy state  $H_2^0$  of an electroweak singlet extension of the Standard Model via the decays of  $H_2^0$  to  $W^+W^-$  and  $ZZ$  in up to 5.1 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 7$  TeV and up to 19.7 fb $^{-1}$  at  $E_{cm} = 8$  TeV in the range  $m_{H_2^0} = 145\text{--}1000$  GeV. See their Figs. 8 and 9 for limits in the parameter space of the model.
- <sup>10</sup> KHACHATRYAN 15BB search for production of a resonance  $H^0$  decaying to  $\gamma\gamma$  in 19.7 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Fig. 7 for limits on cross section times branching ratio for  $m_{H^0} = 150\text{--}850$  GeV.
- <sup>11</sup> KHACHATRYAN 15N search for production of  $A^0$  decaying to  $ZH^0 \rightarrow \ell^+\ell^-b\bar{b}$  in 19.7 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Fig. 3 for limits on cross section times branching ratios for  $m_{A^0} = 225\text{--}600$  GeV.
- <sup>12</sup> KHACHATRYAN 15o search for production of a high-mass narrow resonance  $A^0$  decaying to  $ZH^0 \rightarrow q\bar{q}\tau^+\tau^-$  in 19.7 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Fig. 6 for limits on cross section times branching ratios for  $m_{A^0} = 800\text{--}2500$  GeV.
- <sup>13</sup> AAD 14AP search for a second  $H^0$  state decaying to  $\gamma\gamma$  in addition to the state at about 125 GeV in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Fig. 4 for limits on cross section times branching ratio for  $m_{H^0} = 65\text{--}600$  GeV.
- <sup>14</sup> AAD 14M search for the decay cascade  $H_2^0 \rightarrow H^\pm W^\mp \rightarrow H^0 W^\pm W^\mp, H^0$  decaying to  $b\bar{b}$  in 20.3 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Table III for limits on cross section times branching ratio for  $m_{H_2^0} = 325\text{--}1025$  GeV and  $m_{H^\pm} = 225\text{--}925$  GeV.
- <sup>15</sup> CHATRCHYAN 14G search for a second  $H^0$  state decaying to  $WW^{(*)}$  in addition to the observed signal at about 125 GeV using 4.9 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 7$  TeV and 19.4 fb $^{-1}$  at  $E_{cm} = 8$  TeV. See their Fig. 21 (right) for cross section limits in the mass range 110–600 GeV.
- <sup>16</sup> KHACHATRYAN 14P search for a second  $H^0$  state decaying to  $\gamma\gamma$  in addition to the observed signal at about 125 GeV using 5.1 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 7$  TeV and 19.7 fb $^{-1}$  at  $E_{cm} = 8$  TeV. See their Figs. 27 and 28 for cross section limits in the mass range 110–150 GeV.
- <sup>17</sup> AALTONEN 13P search for production of a heavy Higgs boson  $H^0$  that decays into a charged Higgs boson  $H^\pm$  and a lighter Higgs boson  $H^0$  via the decay chain  $H^0 \rightarrow H^\pm W^\mp, H^\pm \rightarrow W^\pm H^0, H^0 \rightarrow b\bar{b}$  in the final state  $\ell\nu$  plus 4 jets in 8.7 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. See their Fig. 4 for limits on cross section times branching ratio in the  $m_{H^\pm}\text{--}m_{H^0}$  plane for  $m_{H^0} = 126$  GeV.
- <sup>18</sup> CHATRCHYAN 13B search for  $H^0$  production in the decay chain  $H^0 \rightarrow A^0 A^0, A^0 \rightarrow \mu^+\mu^-$  in 5.3 fb $^{-1}$  of  $pp$  collisions at  $E_{cm} = 7$  TeV. See their Fig. 2 for limits on cross section times branching ratio.
- <sup>19</sup> AALTONEN 11P search in 2.7 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV for the decay chain  $t \rightarrow bH^+, H^+ \rightarrow W^+A^0, A^0 \rightarrow \tau^+\tau^-$  with  $m_{A^0}$  between 4 and 9 GeV. See their Fig. 4 for limits on  $B(t \rightarrow bH^+)$  for  $90 < m_{H^+} < 160$  GeV.
- <sup>20</sup> ABBIENDI 10 search for  $e^+e^- \rightarrow ZH^0$  with the decay chain  $H^0 \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 + (\gamma \text{ or } Z^*)$ , when  $\tilde{\chi}_1^0$  and  $\tilde{\chi}_2^0$  are nearly degenerate. For a mass difference of 2 (4) GeV, a lower limit on  $m_{H^0}$  of 108.4 (107.0) GeV (95% CL) is obtained for SM  $ZH^0$  cross section and  $B(H^0 \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0) = 1$ .
- <sup>21</sup> SCHAEEL 10 search for the process  $e^+e^- \rightarrow H^0 Z$  followed by the decay chain  $H^0 \rightarrow A^0 A^0 \rightarrow \tau^+\tau^-\tau^+\tau^-$  with  $Z \rightarrow \ell^+\ell^-, \nu\bar{\nu}$  at  $E_{cm} = 183\text{--}209$  GeV. For a  $H^0 Z Z$  coupling equal to the SM value,  $B(H^0 \rightarrow A^0 A^0) = B(A^0 \rightarrow \tau^+\tau^-) = 1$ , and  $m_{A^0} = 4\text{--}10$  GeV,  $m_{H^0}$  up to 107 GeV is excluded at 95% CL.
- <sup>22</sup> ABAZOV 09V search for  $H^0$  production followed by the decay chain  $H^0 \rightarrow A^0 A^0 \rightarrow \mu^+\mu^-\mu^+\mu^-$  or  $\mu^+\mu^-\tau^+\tau^-$  in 4.2 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. See their Fig. 3 for limits on  $\sigma(H^0) \cdot B(H^0 \rightarrow A^0 A^0)$  for  $m_{A^0} = 3.6\text{--}19$  GeV.
- <sup>23</sup> ABBIENDI 05A search for  $e^+e^- \rightarrow H_1^0 A^0$  in general Type-II two-doublet models, with decays  $H_1^0, A^0 \rightarrow q\bar{q}, gg, \tau^+\tau^-$ , and  $H_1^0 \rightarrow A^0 A^0$ .
- <sup>24</sup> ABBIENDI 04K search for  $e^+e^- \rightarrow H^0 Z$  with  $H^0$  decaying to two jets of any flavor including  $gg$ . The limit is for SM production cross section with  $B(H^0 \rightarrow jj) = 1$ .
- <sup>25</sup> ABDALLAH 04 consider the full combined LEP and LEP2 datasets to set limits on the Higgs coupling to  $W$  or  $Z$  bosons, assuming SM decays of the Higgs. Results in Fig. 26.
- <sup>26</sup> ACHARD 04B search for  $e^+e^- \rightarrow H^0 Z$  with  $H^0$  decaying to  $b\bar{b}, c\bar{c}$ , or  $gg$ . The limit is for SM production cross section with  $B(H^0 \rightarrow jj) = 1$ .
- <sup>27</sup> ACHARD 04F search for  $H^0$  with anomalous coupling to gauge boson pairs in the processes  $e^+e^- \rightarrow H^0 \gamma, e^+e^- H^0, H^0 Z$  with decays  $H^0 \rightarrow f\bar{f}, \gamma\gamma, Z\gamma$ , and  $W^*W$  at  $E_{cm} = 189\text{--}209$  GeV. See paper for limits.
- <sup>28</sup> ABBIENDI 03F search for  $H^0 \rightarrow$  anything in  $e^+e^- \rightarrow H^0 Z$ , using the recoil mass spectrum of  $Z \rightarrow e^+e^-$  or  $\mu^+\mu^-$ . In addition, it searched for  $Z \rightarrow \nu\bar{\nu}$  and  $H^0 \rightarrow e^+e^-$  or photons. Scenarios with large width or continuum  $H^0$  mass distribution are considered. See their Figs. 11–14 for the results.
- <sup>29</sup> ABBIENDI 03G search for  $e^+e^- \rightarrow H_1^0 Z$  followed by  $H_1^0 \rightarrow A^0 A^0, A^0 \rightarrow c\bar{c}, gg$ , or  $\tau^+\tau^-$  in the region  $m_{H_1^0} = 45\text{--}86$  GeV and  $m_{A^0} = 2\text{--}11$  GeV. See their Fig. 7 for the limits.

- <sup>30</sup> Search for associated production of a  $\gamma\gamma$  resonance with a  $Z$  boson, followed by  $Z \rightarrow q\bar{q}, \ell^+\ell^-,$  or  $\nu\bar{\nu}$ , at  $E_{cm} \leq 209$  GeV. The limit is for a  $H^0$  with SM production cross section and  $B(H^0 \rightarrow f\bar{f})=0$  for all fermions  $f$ .
- <sup>31</sup> For  $B(H^0 \rightarrow \gamma\gamma)=1$ ,  $m_{H^0} > 113.1$  GeV is obtained.
- <sup>32</sup> HEISTER 02M search for  $e^+e^- \rightarrow H^0 Z$ , assuming that  $H^0$  decays to  $q\bar{q}, gg$ , or  $\tau^+\tau^-$  only. The limit assumes SM production cross section.
- <sup>33</sup> ABBIENDI 01E search for neutral Higgs bosons in general Type-II two-doublet models, at  $E_{cm} \leq 189$  GeV. In addition to usual final states, the decays  $H_1^0, A^0 \rightarrow q\bar{q}, gg$  are searched for. See their Figs. 15,16 for excluded regions.
- <sup>34</sup> ACCIARRI 00R search for  $e^+e^- \rightarrow H^0 \gamma$  with  $H^0 \rightarrow b\bar{b}, Z\gamma$ , or  $\gamma\gamma$ . See their Fig. 3 for limits on  $\sigma \cdot B$ . Explicit limits within an effective interaction framework are also given, for which the Standard Model Higgs search results are used in addition.
- <sup>35</sup> ACCIARRI 00R search for the two-photon type processes  $e^+e^- \rightarrow e^+e^- H^0$  with  $H^0 \rightarrow b\bar{b}$  or  $\gamma\gamma$ . See their Fig. 4 for limits on  $\Gamma(H^0 \rightarrow \gamma\gamma) \cdot B(H^0 \rightarrow \gamma\gamma \text{ or } b\bar{b})$  for  $m_{H^0}=70\text{--}170$  GeV.
- <sup>36</sup> GONZALEZ-GARCIA 98B use  $D\bar{D}$  limit for  $\gamma\gamma$  events with missing  $E_T$  in  $p\bar{p}$  collisions (ABBOTT 98) to constrain possible  $ZH$  or  $WH$  production followed by unconventional  $H \rightarrow \gamma\gamma$  decay which is induced by higher-dimensional operators. See their Figs. 1 and 2 for limits on the anomalous couplings.
- <sup>37</sup> KRAWCZYK 97 analyse the muon anomalous magnetic moment in a two-doublet Higgs model (with type II Yukawa couplings) assuming no  $H_1^0 Z Z$  coupling and obtain  $m_{H_1^0} \gtrsim 5$  GeV or  $m_{A^0} \gtrsim 5$  GeV for  $\tan\beta > 50$ . Other Higgs bosons are assumed to be much heavier.
- <sup>38</sup> ALEXANDER 96H give  $B(Z \rightarrow H^0 \gamma) \times B(H^0 \rightarrow q\bar{q}) < 1\text{--}4 \times 10^{-5}$  (95%CL) and  $B(Z \rightarrow H^0 \gamma) \times B(H^0 \rightarrow b\bar{b}) < 0.7\text{--}2 \times 10^{-5}$  (95%CL) in the range  $20 < m_{H^0} < 80$  GeV.

## SEARCHES FOR A HIGGS BOSON WITH STANDARD MODEL COUPLINGS

These listings are based on experimental searches for a scalar boson whose couplings to  $W, Z$  and fermions are precisely those of the Higgs boson predicted by the three-generation Standard Model with the minimal Higgs sector.

For a review and a bibliography, see the review on “Status of Higgs Boson Physics.”

### Direct Mass Limits for $H^0$

The mass limits shown below apply to a Higgs boson  $H^0$  with Standard Model couplings whose mass is a priori unknown. These mass limits are compatible with and independent of the observed signal at about 125 GeV. In particular, the symbol  $H^0$  employed below does not in general refer to the observed signal at about 125 GeV.

The cross section times branching ratio limits quoted in the footnotes below are typically given relative to those of a Standard Model Higgs boson of the relevant mass. These limits can be reinterpreted in terms of more general models (e.g. extended Higgs sectors) in which the Higgs couplings to  $W, Z$  and fermions are re-scaled from their Standard Model values.

All data that have been superseded by newer results are marked as “not used” or have been removed from this compilation, and are documented in previous editions of this Review of Particle Physics.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt; 122 and none 128–1000 (CL = 95%)</b>				
none 145–1000	95	<sup>1</sup> KHACHATRY..15AW	CMS	$pp \rightarrow H^0 X$ combined
none 90–102,	95	<sup>2</sup> AALTONEN	13L CDF	$pp \rightarrow H^0 X$ , combined
none 90–109,	95	<sup>3</sup> AALTONEN	13M TEVA	Tevatron combined
149–182				
none 90–101,	95	<sup>4</sup> ABAZOV	13L D0	$p\bar{p} \rightarrow H^0 X$ , combined
157–178				
none 110–121.5,	95	<sup>5</sup> CHATRCHYAN12N	CMS	$pp \rightarrow H^0 X$ combined
128–145				
>114.1,	95	<sup>6</sup> ABDALLAH	04 DLP	$e^+e^- \rightarrow H^0 Z$
>112.7,	95	<sup>6</sup> ABBIENDI	03B OPAL	$e^+e^- \rightarrow H^0 Z$
>114.4,	95	<sup>6,7</sup> HEISTER	03D LEP	$e^+e^- \rightarrow H^0 Z$
>111.5,	95	<sup>6,8</sup> HEISTER	02 ALEP	$e^+e^- \rightarrow H^0 Z$
>112.0,	95	<sup>6</sup> ACHARD	01C L3	$e^+e^- \rightarrow H^0 Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 132–200	95	<sup>9</sup> AAD	15AA ATLS	$pp \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		<sup>10</sup> AAD	15G ATLS	$pp \rightarrow H^0 W/Z X, H^0 \rightarrow b\bar{b}$
		<sup>11</sup> AAD	14AS ATLS	$pp \rightarrow H^0 X, H^0 \rightarrow \mu\mu$
		<sup>12</sup> AAD	14J ATLS	$pp \rightarrow H^0 X, H^0 \rightarrow Z\gamma$
none 114.5–119,	95	<sup>13</sup> CHATRCHYAN14AA	CMS	$pp \rightarrow H^0 X, H^0 \rightarrow 4\ell$
129.5–832				
none 127–600	95	<sup>14</sup> CHATRCHYAN14AI	CMS	$pp \rightarrow H^0 W/Z X, H^0 \rightarrow b\bar{b}$
		<sup>15</sup> CHATRCHYAN14G	CMS	$pp \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		<sup>16</sup> AALTONEN	13B CDF	$p\bar{p} \rightarrow H^0 W/Z X, H^0 \rightarrow b\bar{b}$
		<sup>17</sup> AALTONEN	13C CDF	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow b\bar{b}$
none 149–172	95	<sup>18</sup> AALTONEN	13K CDF	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		<sup>19</sup> ABAZOV	13E D0	$p\bar{p} \rightarrow H^0 X, 4\ell$
		<sup>20</sup> ABAZOV	13F D0	$p\bar{p} \rightarrow H^0 X, \ell\tau jj$
none 159–176	95	<sup>21</sup> ABAZOV	13G D0	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		<sup>22</sup> ABAZOV	13H D0	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$
		<sup>23</sup> ABAZOV	13I D0	$p\bar{p} \rightarrow H^0 X, \ell\nu jj$

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# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

		24	ABAZOV	13J	D0	$p\bar{p} \rightarrow H^0 X$ , leptonic
		25	ABAZOV	13K	D0	$p\bar{p} \rightarrow H^0 Z X$
		26	CHATRCHYAN	13AL	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \tau\tau, WW^{(*)}, ZZ^{(*)}$
		27	CHATRCHYAN	13BK	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow Z\gamma$
none 145–710	95	28	CHATRCHYAN	13Q	CMS	$p\bar{p} \rightarrow H^0 X$ combined
		29	CHATRCHYAN	13X	CMS	$p\bar{p} \rightarrow H^0 t\bar{t} X$
none 113–122, 128–133, 138–149	95	30	CHATRCHYAN	13Y	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$
none 130–164, 170–180	95	31	CHATRCHYAN	13Y	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ^{*}$
none 129–160	95	32	CHATRCHYAN	13Y	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{*}$
none 111–122, 131–559	95	33	AAD	12AI	ATLS	$p\bar{p} \rightarrow H^0 X$ combined
none 133–261	95	34	AAD	12AJ	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		35	AAD	12BU	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \tau^{+}\tau^{-}$
none 319–558	95	36	AAD	12BZ	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ$
none 300–322, 353–410	95	37	AAD	12CA	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ$
		38	AAD	12CN	ATLS	$p\bar{p} \rightarrow H^0 W/Z X, H^0 \rightarrow b\bar{b}$
none 134–156, 182–233, 256–265, 268–415	95	39	AAD	12CO	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW$
		40	AAD	12D	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ^{(*)}$
none 113–115, 134.5–136	95	41	AAD	12G	ATLS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$
		42	AALTONEN	12AK	CDF	$p\bar{p} \rightarrow H^0 t\bar{t} X$
		43	AALTONEN	12AM	CDF	$p\bar{p} \rightarrow H^0 X$ , inclusive $4\ell$
		44	AALTONEN	12AN	CDF	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$
		45	AALTONEN	12J	CDF	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \tau\tau$
		46	AALTONEN	12Q	CDF	$p\bar{p} \rightarrow H^0 Z X, H^0 \rightarrow b\bar{b}$
none 100–106	95	47	AALTONEN	12T	TEVA	$p\bar{p} \rightarrow H^0 W/Z X, H^0 \rightarrow b\bar{b}$
		48	ABAZOV	12K	D0	$p\bar{p} \rightarrow H^0 W/Z X, H^0 \rightarrow b\bar{b}$
		49,50	CHATRCHYAN	12AY	CMS	$p\bar{p} \rightarrow H^0 W X, H^0 Z X$
		51	CHATRCHYAN	12C	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ$
		52	CHATRCHYAN	12D	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ^{(*)}$
none 129–270	95	53	CHATRCHYAN	12E	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		54	CHATRCHYAN	12F	CMS	$p\bar{p} \rightarrow H^0 W X, H^0 Z X$
none 128–132	95	55	CHATRCHYAN	12G	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \gamma\gamma$
none 134–158, 180–305, 340–465	95	56	CHATRCHYAN	12H	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ^{(*)}$
none 270–440	95	57	CHATRCHYAN	12I	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow ZZ$
		58	CHATRCHYAN	12K	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow \tau^{+}\tau^{-}$
		59	ABAZOV	11G	D0	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		60	CHATRCHYAN	11J	CMS	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW$
none 162–166	95	61	AALTONEN	10F	TEVA	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		62	AALTONEN	10M	TEVA	$p\bar{p} \rightarrow gg X \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		63	AALTONEN	09A	CDF	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{(*)}$
		64	ABAZOV	09A	D0	$H^0 \rightarrow \tau^{+}\tau^{-}$
		65	ABAZOV	06	D0	$p\bar{p} \rightarrow H^0 X, H^0 \rightarrow WW^{*}$
		66	ABAZOV	06O	D0	$p\bar{p} \rightarrow H^0 W X, H^0 \rightarrow WW^{*}$

<sup>1</sup> KHACHATRYAN 15AW search for  $H^0$  production in the decays  $H^0 \rightarrow W^{+}W^{-} \rightarrow \ell\nu\ell\nu, \ell\nu q\bar{q}$ , and  $H^0 \rightarrow ZZ \rightarrow 4\ell, \ell\ell\tau\tau, \ell\ell\nu\nu$ , and  $\ell\ell q\bar{q}$  in up to  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and up to  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$  in the range  $m_{H^0} = 145\text{--}1000 \text{ GeV}$ . See their Fig. 7 for limits on cross section times branching ratio.

<sup>2</sup> AALTONEN 13L combine all CDF searches with  $9.45\text{--}10.0 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(0.45\text{--}4.8)$  times the expected Standard Model cross section is given for  $m_{H^0} = 90\text{--}200 \text{ GeV}$  at 95% CL. An excess of events over background is observed with a local significance of  $2.0 \sigma$  at  $m_{H^0} = 125 \text{ GeV}$ . In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 124 and 203 GeV are excluded at 95% CL.

<sup>3</sup> AALTONEN 13M combine all Tevatron data from the CDF and D0 Collaborations. A limit on cross section times branching ratio which corresponds to  $(0.37\text{--}3.1)$  times the expected Standard Model cross section is given for  $m_{H^0} = 90\text{--}200 \text{ GeV}$  at 95% CL. An excess of events over background is observed with a local significance of  $3.0\sigma$  at  $m_{H^0} = 125 \text{ GeV}$ . In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 121 and 225 GeV are excluded at 95% CL.

<sup>4</sup> ABAZOV 13L combine all D0 results with up to  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(0.66\text{--}3.1)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 90\text{--}200 \text{ GeV}$  at 95% CL. An excess of events over background is observed with a local significance of  $1.7\sigma$  at  $m_{H^0} = 125 \text{ GeV}$ . In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 125 and 218 GeV are excluded at 95% CL.

<sup>5</sup> CHATRCHYAN 12N search for  $H^0$  production in the decays  $H \rightarrow \gamma\gamma, ZZ^{*} \rightarrow 4\ell, WW^{*} \rightarrow \ell\nu\ell\nu, \tau\tau$ , and  $b\bar{b}$  in  $4.9\text{--}5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $5.1\text{--}5.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The expected exclusion region for no signal is  $110\text{--}145 \text{ GeV}$  at 99.9% CL. See also CHATRCHYAN 13Y.

<sup>6</sup> Search for  $e^{+}e^{-} \rightarrow H^0 Z$  at  $E_{\text{cm}} \leq 209 \text{ GeV}$  in the final states  $H^0 \rightarrow b\bar{b}$  with  $Z \rightarrow \ell\ell, \nu\bar{\nu}, q\bar{q}, \tau^{+}\tau^{-}$  and  $H^0 \rightarrow \tau^{+}\tau^{-}$  with  $Z \rightarrow q\bar{q}$ .

<sup>7</sup> Combination of the results of all LEP experiments.

<sup>8</sup> A  $3\sigma$  excess of candidate events compatible with  $m_{H^0}$  near  $114 \text{ GeV}$  is observed in the combined channels  $q\bar{q}q\bar{q}, q\bar{q}\ell\bar{\ell}, q\bar{q}\tau^{+}\tau^{-}$ .

<sup>9</sup> AAD 15AA search for  $H^0 \rightarrow WW^{(*)}$  in  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(0.2\text{--}6)$  times the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}200 \text{ GeV}$  at 95% CL.

<sup>10</sup> AAD 15G search for  $WH^0$  and  $ZH^0$  production followed by  $H^0 \rightarrow b\bar{b}$  in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . A limit on the cross section times branching ratio which corresponds to  $(0.8\text{--}2.6)$  times the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}140 \text{ GeV}$  at 95% CL.

<sup>11</sup> AAD 14AS search for  $H^0 \rightarrow \mu^{+}\mu^{-}$  in  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . A limit on the cross section times branching ratio which corresponds to  $(6.5\text{--}16.8)$  times the expected Standard Model cross section is given for  $m_{H^0} = 120\text{--}150 \text{ GeV}$  at 95% CL.

<sup>12</sup> AAD 14J search for  $H^0 \rightarrow Z\gamma \rightarrow \ell\ell\gamma$  in  $4.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(4\text{--}18)$  times the expected Standard Model cross section is given for  $m_{H^0} = 120\text{--}150 \text{ GeV}$  at 95% CL.

<sup>13</sup> CHATRCHYAN 14AA search for  $H^0$  production in the decay mode  $H^0 \rightarrow ZZ^{(*)} \rightarrow 4\ell$  in  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.7 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The expected exclusion region for no signal is  $115\text{--}740 \text{ GeV}$  at the 95% CL. See their Fig. 18 for cross section limits for  $m_{H^0} = 110\text{--}1000 \text{ GeV}$ .

<sup>14</sup> CHATRCHYAN 14AI search for  $WH^0$  and  $ZH^0$  production followed by  $H^0 \rightarrow b\bar{b}$  in up to  $5.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and up to  $18.9 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . A limit on the cross section times branching ratio which corresponds to  $(1\text{--}3)$  times the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}135 \text{ GeV}$  at 95% CL.

<sup>15</sup> CHATRCHYAN 14G search for  $H^0$  production in the decay mode  $H^0 \rightarrow WW^{(*)} \rightarrow \ell\nu\ell\nu$  in  $4.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$  and  $19.4 \text{ fb}^{-1}$  at  $E_{\text{cm}} = 8 \text{ TeV}$ . The expected exclusion region for no signal is  $115\text{--}600 \text{ GeV}$  at the 95% CL. See their Fig. 21 (left) for cross section limits for  $m_{H^0} = 110\text{--}600 \text{ GeV}$ .

<sup>16</sup> AALTONEN 13B search for associated  $H^0 Z$  production in the final state  $H^0 \rightarrow b\bar{b}, Z \rightarrow \nu\bar{\nu}$ , and  $H^0 W$  production in  $H^0 \rightarrow b\bar{b}, W \rightarrow \ell\nu$  ( $\ell$  not identified) with an improved  $b$  identification algorithm in  $9.45 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(0.72\text{--}11.8)$  times the expected Standard Model cross section is given for  $m_{H^0} = 90\text{--}150 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 3.06, where 3.33 is expected for no signal.

<sup>17</sup> AALTONEN 13C search for associated  $H^0 W$  and  $H^0 Z$  as well as vector-boson fusion  $H^0 q\bar{q}$  production in the final state  $H^0 \rightarrow b\bar{b}, W/Z \rightarrow q\bar{q}$  with  $9.45 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which is  $(7.0\text{--}64.6)$  times larger than the expected Standard Model cross section is given in the range  $m_{H^0} = 100\text{--}150 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 9.0, where 11.0 is expected for no signal.

<sup>18</sup> AALTONEN 13K search for  $H^0$  production (with a possible additional  $W$  or  $Z$ ) in the final state  $H^0 \rightarrow WW^{(*)} \rightarrow \ell\nu\ell\nu$  in  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(0.49\text{--}14.1)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 110\text{--}200 \text{ GeV}$  at 95% CL. The limit at  $m_{H^0} = 125 \text{ GeV}$  is 3.26, where 3.25 is expected for no signal. In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 124 and 200 GeV are excluded at 95% CL.

<sup>19</sup> ABAZOV 13E search for  $H^0$  production in four-lepton final states from  $H^0 \rightarrow ZZ^{(*)}$  and  $H^0 Z$  in  $9.6\text{--}9.8 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(8.6\text{--}78.9)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 115\text{--}200 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 42.3, where 42.8 is expected for no signal.

<sup>20</sup> ABAZOV 13F search for  $H^0$  production in final states  $e\tau jj$  and  $\mu\tau jj$  in  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . The search is sensitive to  $H \rightarrow \tau\tau$  and  $H \rightarrow WW^{(*)}$ . A limit on cross section times branching ratio which corresponds to  $(9.4\text{--}17.9)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 105\text{--}150 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 11.3, where 9.0 is expected for no signal.

<sup>21</sup> ABAZOV 13G search for  $H^0$  production in final states  $H^0 \rightarrow WW^{(*)} \rightarrow \ell^{+}\nu\ell^{-}\nu$  in  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$  and give a limit on cross section times branching ratio for  $m_{H^0} = 100\text{--}150 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 4.1, where 3.4 is expected for no signal. In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 125 and 218 GeV are excluded at 95% CL.

<sup>22</sup> ABAZOV 13H search for  $H^0$  production with the decay  $H^0 \rightarrow \gamma\gamma$  in  $9.6 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(8.3\text{--}25.4)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 100\text{--}150 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 12.8, where 8.7 is expected for no signal.

<sup>23</sup> ABAZOV 13I search for  $H^0$  production in the final state with one lepton and two or more jets plus missing  $E_T$  with  $b$  identification in  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . The search is mainly sensitive to  $H^0 W \rightarrow b\bar{b}\ell\nu, H^0 \rightarrow WW^{(*)} \rightarrow \ell\nu q\bar{q}$ , and  $H^0 V \rightarrow V W W^{(*)} \rightarrow \ell\nu q\bar{q} q\bar{q}$  ( $V = W, Z$ ). A limit on cross section times branching ratio which corresponds to  $(1.3\text{--}11.4)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 90\text{--}200 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 5.8, where 4.7 is expected for no signal. In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 150 and 188 GeV are excluded at 95% CL.

<sup>24</sup> ABAZOV 13J search for  $H^0$  production in the final states  $e\mu\mu, e\mu\mu, \mu\tau\tau$ , and  $e^{\pm}\mu^{\pm}$  in  $8.6\text{--}9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . The search is sensitive to  $WH^0, ZH^0$  and gluon fusion production with  $H^0 \rightarrow WW^{(*)}, ZZ^{(*)}$ , decaying to leptonic final states, and to  $WH^0, ZH^0$  production with  $H^0 \rightarrow \tau^{+}\tau^{-}$ . A limit on cross section times branching ratio which corresponds to  $(4.4\text{--}12.7)$  times the expected Standard Model cross section is given in the range  $m_{H^0} = 100\text{--}200 \text{ GeV}$  at 95% CL. The limit for  $m_{H^0} = 125 \text{ GeV}$  is 8.4, where 6.3 is expected for no signal.

<sup>25</sup> ABAZOV 13K search for associated  $H^0 Z$  production in the final states  $\ell\ell b\bar{b}$  with  $b$  identification in  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A limit on cross section times branching ratio which corresponds to  $(1.8\text{--}5.3)$  times the expected Standard Model

# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

- cross section is given for  $m_{H^0} = 90\text{--}150$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 7.1, where 5.1 is expected for no signal.
- 26 CHATRCHYAN 13AL search for  $H^0 \rightarrow \tau^+\tau^-, WW^*$ , and  $ZZ^*$  in 5.1 fb $^{-1}$  and 5.3 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  and 8 TeV. In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 110 and 600 GeV are excluded at 99% CL.
- 27 CHATRCHYAN 13BK search for  $H^0 \rightarrow Z\gamma \rightarrow \ell\ell\gamma$  in 5.0 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 19.6 fb $^{-1}$  at  $E_{\text{cm}} = 8$  TeV. A limit on cross section times branching ratio which corresponds to (4–25) times the expected Standard Model cross section is given in the range  $m_{H^0} = 120\text{--}160$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 9.5, where 10 is expected for no signal.
- 28 CHATRCHYAN 13Q search for  $H^0$  production in the decays  $H^0 \rightarrow W^+W^- \rightarrow \ell\nu\ell\nu$ ,  $\ell\nu q\bar{q}$  and  $H^0 \rightarrow ZZ \rightarrow 4\ell$ ,  $\ell\ell\tau\tau$ ,  $\ell\ell\nu\nu$ , and  $\ell\ell q\bar{q}$  in up to 5.1 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and up to 5.3 fb $^{-1}$  at  $E_{\text{cm}} = 8$  TeV in the range  $m_{H^0} = 145\text{--}1000$  GeV. Superseded by KHACHATRYAN 15AW.
- 29 CHATRCHYAN 13X search for  $H^0 t\bar{t}$  production followed by  $H^0 \rightarrow b\bar{b}$ , one top decaying to  $\ell\nu$  and the other to either  $\ell\nu$  or  $q\bar{q}$  in 5.0 fb $^{-1}$  and 5.1 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  and 8 TeV. A limit on cross section times branching ratio which corresponds to (4.0–8.6) times the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}140$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 5.8, where 5.2 is expected for no signal.
- 30 CHATRCHYAN 13Y search for  $H^0$  production in the decay  $H \rightarrow \gamma\gamma$  in 5.1 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.3 fb $^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The expected exclusion region for no signal is 110–144 GeV at 95% CL.
- 31 CHATRCHYAN 13Y search for  $H^0$  production in the decay  $H \rightarrow ZZ^* \rightarrow 4\ell$  in 5.0 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.3 fb $^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The expected exclusion region for no signal is 120–180 GeV at 95% CL.
- 32 CHATRCHYAN 13Y search for  $H^0$  production in the decay  $H \rightarrow WW^* \rightarrow \ell\nu\ell\nu$  in 4.9 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV and 5.3 fb $^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The expected exclusion region for no signal is 122–160 GeV at 95% CL.
- 33 AAD 12AI search for  $H^0$  production in  $pp$  collisions for the final states  $H^0 \rightarrow ZZ^*$ ,  $\gamma\gamma$ ,  $WW^*$ ,  $b\bar{b}$ ,  $\tau\tau$  with 4.6–4.8 fb $^{-1}$  at  $E_{\text{cm}} = 7$  TeV, and  $H^0 \rightarrow ZZ^* \rightarrow 4\ell$ ,  $\gamma\gamma$ ,  $WW^*$   $\rightarrow e\nu\mu\nu$  with 5.8–5.9 fb $^{-1}$  at  $E_{\text{cm}} = 8$  TeV. The 99% CL excluded range is 113–114, 117–121, and 132–527 GeV. An excess of events over background with a local significance of 5.9  $\sigma$  is observed at  $m_{H^0} = 126$  GeV.
- 34 AAD 12AJ search for  $H^0$  production in the decay  $H^0 \rightarrow WW^* \rightarrow \ell\nu\ell\nu$  with 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which corresponds to (0.2–10) times the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}600$  GeV at 95% CL.
- 35 AAD 12BU search for  $H^0$  production in the decay  $H \rightarrow \tau^+\tau^-$  with 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (2.9–11.7) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 100\text{--}150$  GeV at 95% CL.
- 36 AAD 12BZ search for  $H^0$  production in the decay  $H \rightarrow ZZ \rightarrow \ell^+\ell^-\nu\bar{\nu}$  with 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which corresponds to (0.2–4) times the expected Standard Model cross section is given for  $m_{H^0} = 200\text{--}600$  GeV at 95% CL.
- 37 AAD 12CA search for  $H^0$  production in the decay  $H \rightarrow ZZ \rightarrow \ell^+\ell^-q\bar{q}$  with 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which corresponds to (0.7–9) times the expected Standard Model cross section is given for  $m_{H^0} = 200\text{--}600$  GeV at 95% CL.
- 38 AAD 12CN search for associated  $H^0 W$  and  $H^0 Z$  production in the channels  $W \rightarrow \ell\nu$ ,  $Z \rightarrow \ell^+\ell^-$ ,  $\nu\bar{\nu}$ , and  $H^0 \rightarrow b\bar{b}$ , with 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (2.5–5.5) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}130$  GeV at 95% CL.
- 39 AAD 12CO search for  $H^0$  production in the decay  $H \rightarrow WW \rightarrow \ell\nu q\bar{q}$  with 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (1.9–10) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 300\text{--}600$  GeV at 95% CL.
- 40 AAD 12D search for  $H^0$  production with  $H \rightarrow ZZ^* \rightarrow 4\ell$  in 4.8 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 110\text{--}600$  GeV. An excess of events over background with a local significance of 2.1  $\sigma$  is observed at 125 GeV.
- 41 AAD 12G search for  $H^0$  production with  $H \rightarrow \gamma\gamma$  in 4.9 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 110\text{--}150$  GeV. An excess of events over background with a local significance of 2.8  $\sigma$  is observed at 126.5 GeV.
- 42 AALTONEN 12AK search for associated  $H^0 t\bar{t}$  production in the decay chain  $t\bar{t} \rightarrow W W b\bar{b} \rightarrow \ell\nu q\bar{q} b\bar{b}$  with 9.45 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which is (10–40) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 100\text{--}150$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 20.5, where 12.6 is expected.
- 43 AALTONEN 12AM search for  $H^0$  production in inclusive four-lepton final states coming from  $H^0 \rightarrow ZZ, H^0 Z \rightarrow WW^* \ell\ell$ , or  $H^0 Z \rightarrow \tau\tau\ell\ell$ , with 9.7 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which is (7.2–42.4) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 120\text{--}300$  GeV at 95% CL. The best limit is for  $m_{H^0} = 200$  GeV.
- 44 AALTONEN 12AN search for  $H^0$  production in the decay  $H^0 \rightarrow \gamma\gamma$  with 10 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which is (7.7–21.3) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 100\text{--}150$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 17.0, where 9.9 is expected.
- 45 AALTONEN 12J search for  $H^0$  production in the decay  $H^0 \rightarrow \tau^+\tau^-$  (one leptonic, the other hadronic) with 6.0 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which is (14.6–70.2) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 100\text{--}150$  GeV at 95% CL. The best limit is for  $m_{H^0} = 120$  GeV.
- 46 AALTONEN 12Q search for associated  $H^0 Z$  production in the final state  $H^0 \rightarrow b\bar{b}$ ,  $Z \rightarrow \ell^+\ell^-$  with 9.45 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which corresponds to (1.0–37.5) times the expected Standard Model cross section is given for  $m_{H^0} = 90\text{--}150$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 7.1, where 3.9 is expected. A broad excess of events for  $m_{H^0} > 110$  GeV is observed, with a local significance of 2.4  $\sigma$  at  $m_{H^0} = 135$  GeV.
- 47 AALTONEN 12T combine AALTONEN 12Q, AALTONEN 12R, AALTONEN 12S, ABAZOV 12O, ABAZOV 12P, and ABAZOV 12K. An excess of events over background is observed which is most significant in the region  $m_{H^0} = 120\text{--}135$  GeV, with a local significance of up to 3.3  $\sigma$ . The local significance at  $m_{H^0} = 125$  GeV is 2.8  $\sigma$ , which corresponds to  $(\sigma(H^0 W) + \sigma(H^0 Z)) B(H^0 \rightarrow b\bar{b}) = (0.23^{+0.09}_{-0.06})$  pb, compared to the Standard Model expectation at  $m_{H^0} = 125$  GeV of  $0.12 \pm 0.01$  pb.
- 48 ABAZOV 12K search for associated  $H^0 Z$  production in the final state  $H^0 \rightarrow b\bar{b}$ ,  $Z \rightarrow \nu\bar{\nu}$ , and  $H^0 W$  production with  $W \rightarrow \ell\nu$  ( $\ell$  not identified) with 9.5 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. A limit on cross section times branching ratio which is (1.9–16.8) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 100\text{--}150$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 4.3, where 3.9 is expected.
- 49 CHATRCHYAN 12AY search for associated  $H^0 W$  and  $H^0 Z$  production in the channels  $W \rightarrow \ell\nu$ ,  $Z \rightarrow \ell^+\ell^-$ , and  $H^0 \rightarrow \tau\tau$ ,  $WW^*$ , with 5 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (3.1–9.1) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}200$  GeV at 95% CL.
- 50 CHATRCHYAN 12AY combine CHATRCHYAN 12F and CHATRCHYAN 12AO in addition and give a limit on cross section times branching ratio which is (2.1–3.7) times larger than the expected Standard Model cross section for  $m_{H^0} = 110\text{--}170$  GeV at 95% CL. The limit for  $m_{H^0} = 125$  GeV is 3.3.
- 51 CHATRCHYAN 12C search for  $H^0$  production with  $H \rightarrow ZZ \rightarrow \ell^+\ell^-\tau^+\tau^-$  in 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (4–12) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 190\text{--}600$  GeV at 95% CL. The best limit is at  $m_{H^0} = 200$  GeV.
- 52 CHATRCHYAN 12D search for  $H^0$  production with  $H \rightarrow ZZ^* \rightarrow \ell^+\ell^-q\bar{q}$  in 4.6 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which corresponds to (1–22) times the expected Standard Model cross section is given for  $m_{H^0} = 130\text{--}164$  GeV, 200–600 GeV at 95% CL. The best limit is at  $m_{H^0} = 230$  GeV. In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values in the ranges  $m_{H^0} = 154\text{--}161$  GeV and 200–470 GeV are excluded at 95% CL.
- 53 CHATRCHYAN 12E search for  $H^0$  production with  $H \rightarrow WW^* \rightarrow \ell^+\nu\ell^-\bar{\nu}$  in 4.6 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 110\text{--}600$  GeV.
- 54 CHATRCHYAN 12F search for associated  $H^0 W$  and  $H^0 Z$  production followed by  $W \rightarrow \ell\nu$ ,  $Z \rightarrow \ell^+\ell^-$ ,  $\nu\bar{\nu}$ , and  $H^0 \rightarrow b\bar{b}$ , in 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (3.1–9.0) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}135$  GeV at 95% CL. The best limit is at  $m_{H^0} = 110$  GeV.
- 55 CHATRCHYAN 12G search for  $H^0$  production with  $H \rightarrow \gamma\gamma$  in 4.8 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 110\text{--}150$  GeV. An excess of events over background with a local significance of 3.1  $\sigma$  is observed at 124 GeV.
- 56 CHATRCHYAN 12H search for  $H^0$  production with  $H \rightarrow ZZ^* \rightarrow 4\ell$  in 4.7 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 110\text{--}600$  GeV. Excesses of events over background are observed around 119, 126 and 320 GeV. The region  $m_{H^0} = 114.4\text{--}134$  GeV remains consistent with the expectation for the production of a SM-like Higgs boson.
- 57 CHATRCHYAN 12I search for  $H^0$  production with  $H \rightarrow ZZ \rightarrow \ell^+\ell^-\nu\bar{\nu}$  in 4.6 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV in the mass range  $m_{H^0} = 250\text{--}600$  GeV.
- 58 CHATRCHYAN 12K search for  $H^0$  production in the decay  $H \rightarrow \tau^+\tau^-$  with 4.6 fb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. A limit on cross section times branching ratio which is (3.2–7.0) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 110\text{--}145$  GeV at 95% CL.
- 59 ABAZOV 11G search for  $H^0$  production in 5.4 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV in the decay mode  $H^0 \rightarrow WW^* \rightarrow \ell\nu q\bar{q}$  (and processes with similar final states). A limit on cross section times branching ratio which is (3.9–37) times larger than the expected Standard Model cross section is given for  $m_{H^0} = 115\text{--}200$  GeV at 95% CL. The best limit is at  $m_{H^0} = 160$  GeV.
- 60 CHATRCHYAN 11J search for  $H^0$  production with  $H \rightarrow W^+W^- \rightarrow \ell\ell\nu\nu$  in 36 pb $^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. See their Fig. 6 for a limit on cross section times branching ratio for  $m_{H^0} = 120\text{--}600$  GeV at 95% CL. In the Standard Model with an additional generation of heavy quarks and leptons which receive their masses via the Higgs mechanism,  $m_{H^0}$  values between 144 and 207 GeV are excluded at 95% CL.
- 61 AALTONEN 10F combine searches for  $H^0$  decaying to  $W^+W^-$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with 4.8 fb $^{-1}$  (CDF) and 5.4 fb $^{-1}$  (DØ).
- 62 AALTONEN 10M combine searches for  $H^0$  decaying to  $W^+W^-$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with 4.8 fb $^{-1}$  (CDF) and 5.4 fb $^{-1}$  (DØ) and derive limits  $\sigma(p\bar{p} \rightarrow H^0) \cdot B(H^0 \rightarrow W^+W^-) < (1.75\text{--}0.38)$  pb for  $m_H = 120\text{--}165$  GeV, where  $H^0$  is produced in  $g g$  fusion. In the Standard Model with an additional generation of heavy quarks,  $m_{H^0}$  between 131 and 204 GeV is excluded at 95% CL.
- 63 AALTONEN 09A search for  $H^0$  production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV in the decay mode  $H^0 \rightarrow WW^* \rightarrow \ell^+\ell^-\nu\bar{\nu}$ . A limit on  $\sigma(H^0) \cdot B(H^0 \rightarrow WW^*)$  between 0.7 and 2.5 pb (95% CL) is given for  $m_{H^0} = 110\text{--}200$  GeV, which is 1.7–4.5 times larger than the expected Standard Model cross section. The best limit is obtained for  $m_{H^0} = 160$  GeV.
- 64 ABAZOV 09U search for  $H^0 \rightarrow \tau^+\tau^-$  with  $\tau \rightarrow$  hadrons in 1 fb $^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV. The production mechanisms include associated  $W/Z+H^0$  production, weak boson fusion, and gluon fusion. A limit (95% CL) is given for  $m_{H^0} = 105\text{--}145$  GeV, which is 20–82 times larger than the expected Standard Model cross section. The limit for  $m_{H^0} = 115$  GeV is 29 times larger than the expected Standard Model cross section.
- 65 ABAZOV 06 search for Higgs boson production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with the decay chain  $H^0 \rightarrow WW^* \rightarrow \ell^\pm\nu\ell^\mp\bar{\nu}$ . A limit  $\sigma(H^0) \cdot B(H^0 \rightarrow WW^*) <$

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# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for

(5.6–3.2) pb (95 %CL) is given for  $m_{H^0} = 120\text{--}200$  GeV, which far exceeds the expected Standard Model cross section.  
 66 ABABOV 060 search for associated  $H^0 W$  production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with the decay  $H^0 \rightarrow WW^*$ , in the final states  $\ell^\pm \ell'^\mp \nu \nu' X$  where  $\ell = e, \mu$ .  
 A limit  $\sigma(H^0 W) \cdot \text{B}(H^0 \rightarrow WW^*) < (3.2\text{--}2.8)$  pb (95 %CL) is given for  $m_{H^0} = 115\text{--}175$  GeV, which far exceeds the expected Standard Model cross section.

### Indirect Mass Limits for $H^0$ from Electroweak Analysis

The mass limits shown below apply to a Higgs boson  $H^0$  with Standard Model couplings whose mass is a priori unknown.

For limits obtained before the direct measurement of the top quark mass, see the 1996 (Physical Review **D54** 1 (1996)) Edition of this Review. Other studies based on data available prior to 1996 can be found in the 1998 Edition (The European Physical Journal **C3** 1 (1998)) of this Review.

VALUE (GeV)	DOCUMENT ID	TECN
$94^{+25}_{-22}$	<sup>1</sup> BAAK	12A RVUE
$91^{+30}_{-23}$	<sup>2</sup> BAAK	12 RVUE
$91^{+31}_{-24}$	<sup>3</sup> ERLER	10A RVUE
$129^{+74}_{-49}$	<sup>4</sup> LEP-SLC	06 RVUE

• • • We do not use the following data for averages, fits, limits, etc. • • •

- <sup>1</sup> BAAK 12A make Standard Model fits to Z and neutral current parameters,  $m_t$ ,  $m_W$ , and  $\Gamma_W$  measurements available in 2012 (using also preliminary data). The quoted result is obtained from a fit that does not include the measured mass value of the signal observed at the LHC and also no limits from direct Higgs searches.
- <sup>2</sup> BAAK 12 make Standard Model fits to Z and neutral current parameters,  $m_t$ ,  $m_W$ , and  $\Gamma_W$  measurements available in 2010 (using also preliminary data). The quoted result is obtained from a fit that does not include the limit from the direct Higgs searches. The result including direct search data from LEP2, the Tevatron and the LHC is  $120^{+12}_{-5}$  GeV.
- <sup>3</sup> ERLER 10A makes Standard Model fits to Z and neutral current parameters,  $m_t$ ,  $m_W$  measurements available in 2009 (using also preliminary data). The quoted result is obtained from a fit that does not include the limits from the direct Higgs searches. With direct search data from LEP2 and Tevatron added to the fit, the 90% CL (99% CL) interval is 115–148 (114–197) GeV.
- <sup>4</sup> LEP-SLC 06 make Standard Model fits to Z parameters from LEP/SLC and  $m_t$ ,  $m_W$ , and  $\Gamma_W$  measurements available in 2005 with  $\Delta\alpha_{\text{had}}^{(5)}(m_Z) = 0.02758 \pm 0.00035$ . The 95% CL limit is 285 GeV.

### SEARCHES FOR NEUTRAL HIGGS BOSONS REFERENCES

AAD	16C	JHEP 1601 032	G. Aad et al.	(ATLAS Collab.)
KHACHATRYAN...	16A	PL B752 221	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	16F	JHEP 1601 079	V. Khachatryan et al.	(CMS Collab.)
AAD	15AA	PR D92 012006	G. Aad et al.	(ATLAS Collab.)
AAD	15BD	EPJ C75 337	G. Aad et al.	(ATLAS Collab.)
AAD	15BH	EPJ C75 299	G. Aad et al.	(ATLAS Collab.)
AAD	15BK	EPJ C75 412	G. Aad et al.	(ATLAS Collab.)
AAD	15BZ	PR D92 052002	G. Aad et al.	(ATLAS Collab.)
AAD	15CF	PR D92 092004	G. Aad et al.	(ATLAS Collab.)
AAD	15G	JHEP 1501 069	G. Aad et al.	(ATLAS Collab.)
AAD	15H	PRL 114 081802	G. Aad et al.	(ATLAS Collab.)
AAD	15S	PL B744 163	G. Aad et al.	(ATLAS Collab.)
KHACHATRYAN...	15AW	JHEP 1510 144	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15AY	JHEP 1511 071	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15BB	PL B750 494	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15N	PL B748 221	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	15O	PL B748 255	V. Khachatryan et al.	(CMS Collab.)
LEES	15H	PR D91 071102	J.P. Lees et al.	(BABAR Collab.)
AAD	14AP	PRL 113 171801	G. Aad et al.	(ATLAS Collab.)
AAD	14AS	PL B738 168	G. Aad et al.	(ATLAS Collab.)
AAD	14AW	JHEP 1411 056	G. Aad et al.	(ATLAS Collab.)
AAD	14BA	JHEP 1411 088	G. Aad et al.	(ATLAS Collab.)
AAD	14J	PL B732 8	G. Aad et al.	(ATLAS Collab.)
AAD	14M	PR D89 032002	G. Aad et al.	(ATLAS Collab.)
AAD	14O	PRL 112 201802	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN	14AA	PR D89 092007	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14AI	PR D89 012003	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14B	EPJ C74 2980	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14G	JHEP 1401 096	S. Chatrchyan et al.	(CMS Collab.)
KHACHATRYAN...	14M	JHEP 1410 160	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	14P	EPJ C74 3076	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN...	14Q	PR D90 112013	V. Khachatryan et al.	(CMS Collab.)
AAD	13AG	PL B721 32	G. Aad et al.	(ATLAS Collab.)
AAD	13AT	NJP 15 043009	G. Aad et al.	(ATLAS Collab.)
AAD	13O	JHEP 1302 095	G. Aad et al.	(ATLAS Collab.)
AAD	13T	JHEP 1305 132	R. Aaij et al.	(LHCb Collab.)
AALTONEN	13B	PR D87 052008	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	13C	JHEP 1302 004	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	13K	PR D88 052012	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	13L	PR D88 052013	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	13M	PR D88 052014	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	13P	PRL 110 121801	T. Aaltonen et al.	(CDF and DO Collabs.)
ABAZOV	13E	PR D88 032008	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13F	PR D88 052005	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13G	PR D88 052006	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13H	PR D88 052007	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13I	PR D88 052008	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13J	PR D88 052009	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13K	PR D88 052010	V.M. Abazov et al.	(DO Collab.)
ABAZOV	13L	PR D88 052011	V.M. Abazov et al.	(DO Collab.)
CARENA	13	EPJ C73 2552	M. Carena et al.	(CMS Collab.)
CHATRCHYAN	13AG	PL B722 207	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13AL	PL B725 36	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13BJ	PL B726 564	S. Chatrchyan et al.	(CMS Collab.)

CHATRCHYAN	13BK	PL B726 587	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13X	EPJ C73 2469	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13X	JHEP 1305 145	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	13X	JHEP 1306 081	S. Chatrchyan et al.	(CMS Collab.)
LEES	13C	PR D87 031102	J.P. Lees et al.	(BABAR Collab.)
LEES	13L	PR D88 031701	J.P. Lees et al.	(BABAR Collab.)
LEES	13R	PR D88 071102	J.P. Lees et al.	(BABAR Collab.)
AAD	12AI	PL B716 1	G. Aad et al.	(ATLAS Collab.)
AAD	12AJ	PL B716 62	G. Aad et al.	(ATLAS Collab.)
AAD	12AQ	PRL 108 251801	G. Aad et al.	(ATLAS Collab.)
AAD	12BU	JHEP 1209 070	G. Aad et al.	(ATLAS Collab.)
AAD	12BZ	PL B717 29	G. Aad et al.	(ATLAS Collab.)
AAD	12CA	PL B717 70	G. Aad et al.	(ATLAS Collab.)
AAD	12CN	PL B718 369	G. Aad et al.	(ATLAS Collab.)
AAD	12CO	PL B718 391	G. Aad et al.	(ATLAS Collab.)
AAD	12D	PL B710 383	G. Aad et al.	(ATLAS Collab.)
AAD	12G	PRL 108 111803	G. Aad et al.	(ATLAS Collab.)
AAD	12N	EPJ C72 2157	G. Aad et al.	(ATLAS Collab.)
AALTONEN	12AB	PR D85 092001	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12AK	PRL 109 181802	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12AM	PR D86 072012	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12AN	PL B717 173	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12AQ	PR D86 091101	T. Aaltonen et al.	(CDF and DO Collabs.)
AALTONEN	12J	PRL 108 181804	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12Q	PRL 109 111803	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12R	PRL 109 111804	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12S	PRL 109 111805	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12T	PRL 109 071804	T. Aaltonen et al.	(CDF and DO Collabs.)
AALTONEN	12U	PR D85 012007	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	12V	PR D85 032005	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	12G	PL B710 569	V.M. Abazov et al.	(DO Collab.)
ABAZOV	12K	PL B716 285	V.M. Abazov et al.	(DO Collab.)
ABAZOV	12O	PRL 109 121803	V.M. Abazov et al.	(DO Collab.)
ABAZOV	12P	PRL 109 121804	V.M. Abazov et al.	(DO Collab.)
BAAK	12	EPJ C72 2003	M. Baak et al.	(Glitter Group)
BAAK	12A	EPJ C72 2205	M. Baak et al.	(Glitter Group)
CHATRCHYAN	12AO	JHEP 1209 111	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12AY	JHEP 1211 088	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12C	JHEP 1203 081	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12D	JHEP 1204 036	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12E	PL B710 91	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12F	PL B710 284	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12G	PL B710 403	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12H	PRL 108 111804	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12I	JHEP 1203 040	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12K	PL B713 68	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12N	PL B716 30	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12P	PRL 109 121801	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN	11V	PRL 107 031801	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	11G	PRL 106 171802	V.M. Abazov et al.	(DO Collab.)
ABAZOV	11K	PL B698 97	V.M. Abazov et al.	(DO Collab.)
ABAZOV	11W	PRL 107 121801	V.M. Abazov et al.	(DO Collab.)
ABOUZAID	11A	PRL 107 201803	E. Abouzaid et al.	(KTEV Collab.)
CHATRCHYAN	11J	PL B699 25	S. Chatrchyan et al.	(CMS Collab.)
DEL-AMO-SA...	11J	PRL 107 021804	P. del Amo Sanchez et al.	(BABAR Collab.)
LEES	11H	PRL 107 221803	J.P. Lees et al.	(BABAR Collab.)
AALTONEN	10F	PRL 104 061802	T. Aaltonen et al.	(CDF and DO Collabs.)
AALTONEN	10M	PR D82 011102	T. Aaltonen et al.	(CDF and DO Collabs.)
ABBIENDI	10	PL B682 381	G. Abbiendi et al.	(OPAL Collab.)
ANDREAS	10	JHEP 1008 003	S. Andreas et al.	(DESY)
ERLER	10A	PR D81 051301	J. Erler	(UNAM)
HYUN	10	PRL 105 091801	H.J. Hyun et al.	(BELLE Collab.)
SCHAE	10	JHEP 1005 049	S. Schae et al.	(ALPHI Collab.)
AALTONEN	09A	PRL 102 021802	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	09AB	PRL 103 061803	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	09AR	PRL 103 201801	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	09U	PRL 102 251801	V.M. Abazov et al.	(DO Collab.)
ABAZOV	09V	PRL 103 061801	V.M. Abazov et al.	(DO Collab.)
AUBERT	09P	PRL 103 181801	B. Aubert et al.	(BABAR Collab.)
AUBERT	09Z	PRL 103 081803	B. Aubert et al.	(BABAR Collab.)
TUNG	09	PRL 102 051802	Y.C. Tung et al.	(KEK E391a Collab.)
ABAZOV	08U	PRL 101 051801	V.M. Abazov et al.	(DO Collab.)
ABDALLAH	08B	EPJ C54 1	J. Abdallah et al.	(DELPHI Collab.)
Also		EPJ C56 165 (errat.)	J. Abdallah et al.	(DELPHI Collab.)
LOVE	08	PRL 101 151802	W. Love et al.	(CDF Collab.)
ABBIENDI	07	EPJ C49 457	G. Abbiendi et al.	(OPAL Collab.)
BESSON	07	PRL 98 052002	D. Besson et al.	(CLEO Collab.)
SCHAE	07	EPJ C49 439	S. Schae et al.	(ALPHI Collab.)
ABAZOV	06	PRL 96 011801	V.M. Abazov et al.	(DO Collab.)
ABAZOV	06O	PRL 97 151804	V.M. Abazov et al.	(DO Collab.)
LEP-SLC	06	PRPL 427 257	ALEPH, DELPHI, L3, OPAL, SLD and working groups	
SCHAE	06B	EPJ C47 547	S. Schae et al.	(LEP Collabs.)
ABBIENDI	05A	EPJ C40 317	G. Abbiendi et al.	(OPAL Collab.)
ABDALLAH	05D	EPJ C44 147	J. Abdallah et al.	(DELPHI Collab.)
ACHARD	05	PL B609 35	P. Achard et al.	(L3 Collab.)
ACOSTA	05Q	PR D72 072004	D. Acosta et al.	(CDF Collab.)
PARK	05	PRL 94 021801	H.K. Park et al.	(FNAL HyperCP Collab.)
ABBIENDI	04K	PL B597 11	G. Abbiendi et al.	(OPAL Collab.)
ABDALLAH	04	EPJ C32 145	J. Abdallah et al.	(DELPHI Collab.)
ABDALLAH	04B	EPJ C32 475	J. Abdallah et al.	(DELPHI Collab.)
ABDALLAH	04L	EPJ C35 313	J. Abdallah et al.	(DELPHI Collab.)
ABDALLAH	04O	EPJ C38 1	J. Abdallah et al.	(DELPHI Collab.)
ACHARD	04B	PL B583 14	P. Achard et al.	(L3 Collab.)
ACHARD	04F	PL B589 89	P. Achard et al.	(L3 Collab.)
ABBIENDI	03B	EPJ C26 479	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI	03F	EPJ C27 311	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI	03G	EPJ C27 483	G. Abbiendi et al.	(OPAL Collab.)
ACHARD	03C	PL B568 191	P. Achard et al.	(L3 Collab.)
CARENA	03	EPJ C26 601	M.S. Carena et al.	(ALEPH, DELPHI, L3+)
HEISTER	03D	PL B565 61	A. Heister et al.	(ALEPH, DELPHI, L3+)
Also		ALEPH, DELPHI, L3, OPAL, LEP Higgs Working Group		
ABBIENDI	02D	EPJ C23 397	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI	02F	PL B544 44	G. Abbiendi et al.	(OPAL Collab.)
ACHARD	02C	PL B534 28	P. Achard et al.	(L3 Collab.)
ACHARD	02H	PL B545 30	P. Achard et al.	(L3 Collab.)
AKEROYD	02	PR D66 037702	A.G. Akeroyd et al.	(ALEPH Collab.)
HEISTER	02	PL B526 191	A. Heister et al.	(ALEPH Collab.)
HEISTER	02L	PL B544 16	A. Heister et al.	(ALEPH Collab.)
HEISTER	02M	PL B544 25	A. Heister et al.	(ALEPH Collab.)
ABBIENDI	01E	EPJ C18 425	G. Abbiendi et al.	(OPAL Collab.)
ABREU	01F	PL B507 89	P. Abreu et al.	(DELPHI Collab.)
ACHARD	01C	PL B517 319	P. Achard et al.	(L3 Collab.)
AFFOLDER	01H	PR D64 092002	T. Affolder et al.	(CDF Collab.)
BARATE	01C	PL B499 53	R. Barate et al.	(ALEPH Collab.)
ACCIARRI	00M	PL B485 85	M. Acciarri et al.	(L3 Collab.)
ACCIARRI	00R	PL B489 102	M. Acciarri et al.	(L3 Collab.)
ACCIARRI	00S	PL B489 115	M. Acciarri et al.	(L3 Collab.)
BARATE	00L	PL B487 241	R. Barate et al.	(ALEPH Collab.)
ABBIENDI	99E	EPJ C7 407	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI	99O	PL B464 311	G. Abbiendi et al.	(OPAL Collab.)
ABBOTT	99B	PRL 82 2244	B. Abbott et al.	(DO Collab.)

# Gauge & Higgs Boson Particle Listings

## Neutral Higgs Bosons, Searches for, Charged Higgs Bosons ( $H^\pm$ and $H^{\pm\pm}$ ), Searches for

ABREU	99P	PL B458 431	P. Abreu <i>et al.</i>	(DELPHI Collab.)
CARENA	99B	hep-ph/9912223	M.S. Carena <i>et al.</i>	
CERN-TH/99-374				
ABBOTT	98	PRL 80 442	B. Abbott <i>et al.</i>	(D0 Collab.)
ACKERSTAFF	98S	EPJ C5 19	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	98Y	PL B437 218	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
GONZALEZ-G.	98B	PR D57 7045	M.C. Gonzalez-Garcia, S.M. Lletti, S.F. Novae	
PDG	98	EPJ C3 1	C. Caso <i>et al.</i>	(PDG Collab.)
KRAWCZYK	97	PR D55 6968	M. Krawczyk, J. Zochowski	(WARS)
ALEXANDER	96H	ZPHY C71 1	G. Alexander <i>et al.</i>	(OPAL Collab.)
PDG	96	PR D54 1	R. M. Barnett <i>et al.</i>	(PDG Collab.)
ABREU	95H	ZPHY C67 69	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BALEST	95	PR D51 2053	R. Balest <i>et al.</i>	(CLEO Collab.)
PICH	92	NP B388 31	A. Pich, J. Prades, P. Yepes	(CERN, CPPM)
ANTREASYAN	90C	PL B251 204	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)

## Charged Higgs Bosons ( $H^\pm$ and $H^{\pm\pm}$ ), Searches for

### CONTENTS:

$H^\pm$  (Charged Higgs) Mass Limits

Mass limits for  $H^{\pm\pm}$  (doubly-charged Higgs boson)

– Limits for  $H^{\pm\pm}$  with  $T_3 = \pm 1$

– Limits for  $H^{\pm\pm}$  with  $T_3 = 0$

### $H^\pm$ (Charged Higgs) MASS LIMITS

Unless otherwise stated, the limits below assume  $B(H^\pm \rightarrow \tau^+ \nu) + B(H^\pm \rightarrow c\bar{s}) = 1$ , and hold for all values of  $B(H^\pm \rightarrow \tau^+ \nu_\tau)$ , and assume  $H^\pm$  weak isospin of  $T_3 = \pm 1/2$ . In the following,  $\tan\beta$  is the ratio of the two vacuum expectation values in two-doublet models (2HDM).

The limits are also applicable to point-like technipions. For a discussion of techniparticles, see the Review of Dynamical Electroweak Symmetry Breaking in this Review.

For limits obtained in hadronic collisions before the observation of the top quark, and based on the top mass values inconsistent with the current measurements, see the 1996 (Physical Review **D54** 1 (1996)) Edition of this Review.

Searches in  $e^+e^-$  collisions at and above the  $Z$  pole have conclusively ruled out the existence of a charged Higgs in the region  $m_{H^\pm} \lesssim 45$  GeV, and are meanwhile superseded by the searches in higher energy  $e^+e^-$  collisions at LEP. Results that are by now obsolete are therefore not included in this compilation, and can be found in a previous Edition (The European Physical Journal **C15** 1 (2000)) of this Review.

In the following, and unless otherwise stated, results from the LEP experiments (ALEPH, DELPHI, L3, and OPAL) are assumed to derive from the study of the  $e^+e^- \rightarrow H^\pm H^\mp$  process. Limits from  $b \rightarrow s\gamma$  decays are usually stronger in generic 2HDM models than in Supersymmetric models.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 80	95	1 LEP	13 LEP	$e^+e^- \rightarrow H^\pm H^\mp, E_{cm} \leq 209\text{ GeV}$
> 76.3	95	2 ABBIENDI	12 OPAL	$e^+e^- \rightarrow H^\pm H^\mp, E_{cm} \leq 209\text{ GeV}$
> 74.4	95	ABDALLAH	04i DLPH	$E_{cm} \leq 209\text{ GeV}$
> 76.5	95	ACHARD	03e L3	$E_{cm} \leq 209\text{ GeV}$
> 79.3	95	HEISTER	02P ALEP	$E_{cm} \leq 209\text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		3 AAD	15AF ATLS	$t \rightarrow bH^\pm$
		4 AAD	15AF ATLS	$tH^\pm$
		5 AAD	15M ATLS	$H^\pm \rightarrow W^\pm Z$
		6 KHACHATRYAN..15AX CMS		$t \rightarrow bH^\pm, H^\pm \rightarrow \tau^\pm \nu$
		7 KHACHATRYAN..15AX CMS		$tH^\pm, H^\pm \rightarrow t\bar{b}$
		8 KHACHATRYAN..15AX CMS		$tH^\pm, H^\pm \rightarrow \tau^\pm \nu$
		9 KHACHATRYAN..15BF CMS		$t \rightarrow bH^\pm, H^\pm \rightarrow c\bar{s}$
		10 AAD	14M ATLS	$H_2^0 \rightarrow H^\pm W^\mp \rightarrow H^0 W^\pm W^\mp, H^0 \rightarrow b\bar{b}$
		11 AALTONEN	14A CDF	$t \rightarrow b\tau\nu$
		12 AAD	13AC ATLS	$t \rightarrow bH^\pm$
		13 AAD	13V ATLS	$t \rightarrow bH^\pm, \text{lepton non-universality}$
		14 AAD	12BH ATLS	$t \rightarrow bH^\pm$
		15 CHATRCHYAN 12AA CMS		$t \rightarrow bH^\pm$
		16 AALTONEN	11P CDF	$t \rightarrow bH^\pm, H^\pm \rightarrow W^\pm A^0$
		17 DESCHAMPS	10 RVUE	Type II, flavor physics data
		18 AALTONEN	09AJ CDF	$t \rightarrow bH^\pm$
		19 ABAZOV	09AC D0	$t \rightarrow bH^\pm$
		20 ABAZOV	09AG D0	$t \rightarrow bH^\pm$

		21 ABAZOV	09AI D0	$t \rightarrow bH^\pm$
		22 ABAZOV	09P D0	$H^\pm \rightarrow t\bar{b}$
		23 ABULENCIA	06E CDF	$t \rightarrow bH^\pm$
> 92.0	95	ABBIENDI	04 OPAL	$B(\tau\nu) = 1$
> 76.7	95	24 ABDALLAH	04i DLPH	Type I
		25 ABBIENDI	03 OPAL	$\tau \rightarrow \mu\nu, e\nu$
		26 ABAZOV	02B D0	$t \rightarrow bH^\pm, H \rightarrow \tau\nu$
		27 BORZUMATI	02 RVUE	
		28 ABBIENDI	01Q OPAL	$B \rightarrow \tau\nu_\tau X$
		29 BARATE	01E ALEP	$B \rightarrow \tau\nu_\tau$
		30 GAMBINO	01 RVUE	$b \rightarrow s\gamma$
>315	99	31 AFFOLDER	00i CDF	$t \rightarrow bH^\pm, H \rightarrow \tau\nu$
		ABBIENDI	99E OPAL	$E_{cm} \leq 183\text{ GeV}$
> 59.5	95	32 ABBOTT	99E D0	$t \rightarrow bH^\pm$
		33 ACKERSTAFF	99D OPAL	$\tau \rightarrow e\nu\nu, \mu\nu\nu$
		34 ACCIARRI	97F L3	$B \rightarrow \tau\nu_\tau$
		35 AMMAR	97B CLEO	$\tau \rightarrow \mu\nu$
		36 COARASA	97 RVUE	$B \rightarrow \tau\nu_\tau X$
		37 GUCHAIT	97 RVUE	$t \rightarrow bH^\pm, H \rightarrow \tau\nu$
		38 MANGANO	97 RVUE	$B_u(c) \rightarrow \tau\nu_\tau$
		39 STAHL	97 RVUE	$\tau \rightarrow \mu\nu$
>244	95	40 ALAM	95 CLE2	$b \rightarrow s\gamma$
		41 BUSKULIC	95 ALEP	$b \rightarrow \tau\nu_\tau X$

<sup>1</sup> LEP 13 give a limit that refers to the Type II scenario. The limit for  $B(H^\pm \rightarrow \tau\nu) = 1$  is 94 GeV (95% CL), and for  $B(H^\pm \rightarrow c\bar{s}) = 1$  the region below 80.5 as well as the region 83–88 GeV is excluded (95% CL). LEP 13 also search for the decay mode  $H^\pm \rightarrow A^0 W^*$  with  $A^0 \rightarrow b\bar{b}$ , which is not negligible in Type I models. The limit in Type I models is 72.5 GeV (95% CL) if  $m_{A^0} > 12$  GeV.

<sup>2</sup> ABBIENDI 12 also search for the decay mode  $H^\pm \rightarrow A^0 W^*$  with  $A^0 \rightarrow b\bar{b}$ .

<sup>3</sup> AAD 15AF search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow \tau^\pm \nu$  in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. Upper limits on  $B(t \rightarrow bH^\pm) B(H^\pm \rightarrow \tau\nu)$  between  $2.3 \times 10^{-3}$  and  $1.3 \times 10^{-2}$  (95% CL) are given for  $m_{H^\pm} = 80$ –160 GeV. See their Fig. 8 for the excluded regions in different benchmark scenarios of the MSSM. The region  $m_{H^\pm} < 140$  GeV is excluded for  $\tan\beta > 1$  in the considered scenarios.

<sup>4</sup> AAD 15AF search for  $tH^\pm$  associated production followed by  $H^\pm \rightarrow \tau^\pm \nu$  in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. Upper limits on  $\sigma(tH^\pm) B(H^\pm \rightarrow \tau\nu)$  between 760 and 4.5 fb (95% CL) are given for  $m_{H^\pm} = 180$ –1000 GeV. See their Fig. 8 for the excluded regions in different benchmark scenarios of the MSSM.

<sup>5</sup> AAD 15M search for vector boson fusion production of  $H^\pm$  decaying to  $W^\pm Z \rightarrow q\bar{q}l^+l^-$  in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Fig. 2 for limits on cross section times branching ratio for  $m_{H^\pm} = 200$ –1000 GeV, and Fig. 3 for limits on the triplet vacuum expectation value fraction in the Georgi-Machacek model.

<sup>6</sup> KHACHATRYAN 15AX search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow \tau^\pm \nu$  in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. Upper limits on  $B(t \rightarrow bH^\pm) B(H^\pm \rightarrow \tau\nu)$  between  $1.2 \times 10^{-2}$  and  $1.5 \times 10^{-3}$  (95% CL) are given for  $m_{H^\pm} = 80$ –160 GeV. See their Fig. 11 for the excluded regions in different benchmark scenarios of the MSSM. The region  $m_{H^\pm} < 155$  GeV is excluded for  $\tan\beta > 1$  in the considered scenarios.

<sup>7</sup> KHACHATRYAN 15AX search for  $tH^\pm$  associated production followed by  $H^\pm \rightarrow t\bar{b}$  in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. Upper limits on  $\sigma(tH^\pm) B(H^\pm \rightarrow t\bar{b})$  between 2.0 and 0.13 pb (95% CL) are given for  $m_{H^\pm} = 180$ –600 GeV. See their Fig. 11 for the excluded regions in different benchmark scenarios of the MSSM.

<sup>8</sup> KHACHATRYAN 15AX search for  $tH^\pm$  associated production followed by  $H^\pm \rightarrow \tau^\pm \nu$  in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. Upper limits on  $\sigma(tH^\pm) B(H^\pm \rightarrow \tau\nu)$  between 380 and 25 fb (95% CL) are given for  $m_{H^\pm} = 180$ –600 GeV. See their Fig. 11 for the excluded regions in different benchmark scenarios of the MSSM.

<sup>9</sup> KHACHATRYAN 15BF search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow c\bar{s}$  in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. Upper limits on  $B(t \rightarrow bH^\pm) B(H^\pm \rightarrow c\bar{s})$  between  $1.2 \times 10^{-2}$  and  $6.5 \times 10^{-2}$  (95% CL) are given for  $m_{H^\pm} = 90$ –160 GeV.

<sup>10</sup> AAD 14M search for the decay cascade  $H_2^0 \rightarrow H^\pm W^\mp \rightarrow H^0 W^\pm W^\mp, H^0$  decaying to  $b\bar{b}$  in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 8$  TeV. See their Table III for limits on cross section times branching ratio for  $m_{H_2^0} = 325$ –1025 GeV and  $m_{H^\pm} = 225$ –925 GeV.

<sup>11</sup> AALTONEN 14A measure  $B(t \rightarrow b\tau\nu) = 0.096 \pm 0.028$  using 9 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. For  $m_{H^\pm} = 80$ –140 GeV, this measured value is translated to a limit  $B(t \rightarrow bH^\pm) < 0.059$  at 95% CL assuming  $B(H^\pm \rightarrow \tau^\pm \nu) = 1$ .

<sup>12</sup> AAD 13AC search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow c\bar{s}$  (flavor unidentified) in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. Upper limits on  $B(t \rightarrow bH^\pm)$  between 0.05 and 0.01 (95%CL) are given for  $m_{H^\pm} = 90$ –150 GeV and  $B(H^\pm \rightarrow c\bar{s}) = 1$ .

<sup>13</sup> AAD 13V search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow \tau^\pm \nu$  through violation of lepton universality with 4.6 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. Upper limits on  $B(t \rightarrow bH^\pm)$  between 0.032 and 0.044 (95% CL) are given for  $m_{H^\pm} = 90$ –140 GeV and  $B(H^\pm \rightarrow \tau^\pm \nu) = 1$ . By combining with AAD 12BH, the limits improve to 0.008 to 0.034 for  $m_{H^\pm} = 90$ –160 GeV. See their Fig. 7 for the excluded region in the  $m_h^{\max}$  scenario of the MSSM.

<sup>14</sup> AAD 12BH search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow \tau^\pm \nu$  with 4.6 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. Upper limits on  $B(t \rightarrow bH^\pm)$  between 0.01 and 0.05 (95% CL) are given for  $m_{H^\pm} = 90$ –160 GeV and  $B(H^\pm \rightarrow \tau^\pm \nu) = 1$ . See their Fig. 8 for the excluded region in the  $m_h^{\max}$  scenario of the MSSM.

<sup>15</sup> CHATRCHYAN 12AA search for  $t\bar{t}$  production followed by  $t \rightarrow bH^\pm, H^\pm \rightarrow \tau^\pm \nu$  with 2 fb<sup>-1</sup> of  $pp$  collisions at  $E_{cm} = 7$  TeV. Upper limits on  $B(t \rightarrow bH^\pm)$  between 0.019 and 0.041 (95% CL) are given for  $m_{H^\pm} = 80$ –160 GeV and  $B(H^\pm \rightarrow \tau^\pm \nu) = 1$ .

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# Gauge & Higgs Boson Particle Listings

## Charged Higgs Bosons ( $H^\pm$ and $H^{\pm\pm}$ ), Searches for

- <sup>16</sup> AALTONEN 11P search in  $2.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$  for the decay chain  $t \rightarrow bH^+, H^+ \rightarrow W^+A^0, A^0 \rightarrow \tau^+\tau^-$  with  $m_{A^0}$  between 4 and 9 GeV. See their Fig. 4 for limits on  $B(t \rightarrow bH^+)$  for  $90 < m_{H^+} < 160 \text{ GeV}$ .
- <sup>17</sup> DESCHAMPS 10 make Type II two Higgs doublet model fits to weak leptonic and semileptonic decays,  $b \rightarrow s\gamma, B, B_s$  mixings, and  $Z \rightarrow b\bar{b}$ . The limit holds irrespective of  $\tan\beta$ .
- <sup>18</sup> AALTONEN 09AJ search for  $t \rightarrow bH^+, H^+ \rightarrow c\bar{s}$  in  $t\bar{t}$  events in  $2.2 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . Upper limits on  $B(t \rightarrow bH^+)$  between 0.08 and 0.32 (95% CL) are given for  $m_{H^+} = 60\text{--}150 \text{ GeV}$  and  $B(H^+ \rightarrow c\bar{s}) = 1$ .
- <sup>19</sup> ABAZOV 09AC search for  $t \rightarrow bH^+, H^+ \rightarrow \tau^+\nu$  in  $t\bar{t}$  events in  $0.9 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . Upper limits on  $B(t \rightarrow bH^+)$  between 0.19 and 0.25 (95% CL) are given for  $m_{H^+} = 80\text{--}155 \text{ GeV}$  and  $B(H^+ \rightarrow \tau^+\nu) = 1$ . See their Fig. 4 for an excluded region in a MSSM scenario.
- <sup>20</sup> ABAZOV 09AG measure  $t\bar{t}$  cross sections in final states with  $\ell$  + jets ( $\ell = e, \mu$ ),  $\ell\bar{\ell}$ , and  $\tau\ell$  in  $1 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ , which constrains possible  $t \rightarrow bH^+$  branching fractions. Upper limits (95% CL) on  $B(t \rightarrow bH^+)$  between 0.15 and 0.40 (0.48 and 0.57) are given for  $B(H^+ \rightarrow \tau^+\nu) = 1$  ( $B(H^+ \rightarrow c\bar{s}) = 1$ ) for  $m_{H^+} = 80\text{--}155 \text{ GeV}$ .
- <sup>21</sup> ABAZOV 09AI search for  $t \rightarrow bH^+$  in  $t\bar{t}$  events in  $1 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . Final states with  $\ell$  + jets ( $\ell = e, \mu$ ),  $\ell\bar{\ell}$ , and  $\tau\ell$  are examined. Upper limits on  $B(t \rightarrow bH^+)$  (95% CL) between 0.15 and 0.19 (0.19 and 0.22) are given for  $B(H^+ \rightarrow \tau^+\nu) = 1$  ( $B(H^+ \rightarrow c\bar{s}) = 1$ ) for  $m_{H^+} = 80\text{--}155 \text{ GeV}$ . For  $B(H^+ \rightarrow \tau^+\nu) = 1$  also a simultaneous extraction of  $B(t \rightarrow bH^+)$  and the  $t\bar{t}$  cross section is performed, yielding a limit on  $B(t \rightarrow bH^+)$  between 0.12 and 0.26 for  $m_{H^+} = 80\text{--}155 \text{ GeV}$ . See their Figs. 5–8 for excluded regions in several MSSM scenarios.
- <sup>22</sup> ABAZOV 09P search for  $H^\pm$  production by  $q\bar{q}$  annihilation followed by  $H^\pm \rightarrow t\bar{b}$  decay in  $0.9 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . Cross section limits in several two-doublet models are given for  $m_{H^\pm} = 180\text{--}300 \text{ GeV}$ . A region with  $20 \lesssim \tan\beta \lesssim 70$  is excluded (95% CL) for  $180 \text{ GeV} \lesssim m_{H^\pm} \lesssim 184 \text{ GeV}$  in type-I models.
- <sup>23</sup> ABULECNA 06E search for associated  $H^0 W$  production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96 \text{ TeV}$ . A fit is made for  $t\bar{t}$  production processes in dilepton, lepton + jets, and lepton +  $\tau$  final states, with the decays  $t \rightarrow W^+b$  and  $t \rightarrow H^+b$  followed by  $H^\pm \rightarrow \tau^+\nu, c\bar{s}, t^*\bar{b}$ , or  $W^+H^0$ . Within the MSSM the search is sensitive to the region  $\tan\beta < 1$  or  $> 30$  in the mass range  $m_{H^\pm} = 80\text{--}160 \text{ GeV}$ . See Fig. 2 for the excluded region in a certain MSSM scenario.
- <sup>24</sup> ABDALLAH 04I search for  $e^+e^- \rightarrow H^+H^-$  with  $H^\pm$  decaying to  $\tau\nu, c\bar{s}$ , or  $W^*A^0$  in Type-I two-Higgs-doublet models.
- <sup>25</sup> ABBIENDI 03 give a limit  $m_{H^\pm} > 1.28\tan\beta \text{ GeV}$  (95%CL) in Type II two-doublet models.
- <sup>26</sup> ABAZOV 02B search for a charged Higgs boson in top decays with  $H^\pm \rightarrow \tau^+\nu$  at  $E_{\text{cm}} = 1.8 \text{ TeV}$ . For  $m_{H^\pm} = 75 \text{ GeV}$ , the region  $\tan\beta > 32.0$  is excluded at 95% CL. The excluded mass region extends to over 140 GeV for  $\tan\beta$  values above 100.
- <sup>27</sup> BORZUMATI 02 point out that the decay modes such as  $b\bar{b}W, A^0 W$ , and supersymmetric ones can have substantial branching fractions in the mass range explored at LEP II and Tevatron.
- <sup>28</sup> ABBIENDI 03Q give a limit  $\tan\beta/m_{H^\pm} < 0.53 \text{ GeV}^{-1}$  (95%CL) in Type II two-doublet models.
- <sup>29</sup> BARATE 01E give a limit  $\tan\beta/m_{H^\pm} < 0.40 \text{ GeV}^{-1}$  (90% CL) in Type II two-doublet models. An independent measurement of  $B \rightarrow \tau\nu_\tau X$  gives  $\tan\beta/m_{H^\pm} < 0.49 \text{ GeV}^{-1}$  (90% CL).
- <sup>30</sup> GAMBINO 01 use the world average data in the summer of 2001  $B(b \rightarrow s\gamma) = (3.23 \pm 0.42) \times 10^{-4}$ . The limit applies for Type-II two-doublet models.
- <sup>31</sup> AFFOLDER 00I search for a charged Higgs boson in top decays with  $H^\pm \rightarrow \tau^+\nu$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.8 \text{ TeV}$ . The excluded mass region extends to over 120 GeV for  $\tan\beta$  values above 100 and  $B(\tau\nu) = 1$ . If  $B(t \rightarrow bH^+) \gtrsim 0.6$ ,  $m_{H^\pm}$  up to 160 GeV is excluded. Updates ABE 97L.
- <sup>32</sup> ABBOTT 99C search for a charged Higgs boson in top decays in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.8 \text{ TeV}$ , by comparing the observed  $t\bar{t}$  cross section (extracted from the data assuming the dominant decay  $t \rightarrow bW^+$ ) with theoretical expectation. The search is sensitive to regions of the domains  $\tan\beta \lesssim 1, 50 < m_{H^\pm}(\text{GeV}) \lesssim 120$  and  $\tan\beta \gtrsim 40, 50 < m_{H^\pm}(\text{GeV}) \lesssim 160$ . See Fig. 3 for the details of the excluded region.
- <sup>33</sup> ACKERSTAFF 99D measure the Michel parameters  $\rho, \xi, \eta$ , and  $\xi\delta$  in leptonic  $\tau$  decays from  $Z \rightarrow \tau\tau$ . Assuming  $e\text{--}\mu$  universality, the limit  $m_{H^\pm} > 0.97 \tan\beta \text{ GeV}$  (95% CL) is obtained for two-doublet models in which only one doublet couples to leptons.
- <sup>34</sup> ACCIARRI 97F give a limit  $m_{H^\pm} > 2.6 \tan\beta \text{ GeV}$  (90% CL) from their limit on the exclusive  $B \rightarrow \tau\nu_\tau$  branching ratio.
- <sup>35</sup> AMMAR 97B measure the Michel parameter  $\rho$  from  $\tau \rightarrow e\nu\nu$  decays and assumes  $e/\mu$  universality to extract the Michel  $\eta$  parameter from  $\tau \rightarrow \mu\nu\nu$  decays. The measurement is translated to a lower limit on  $m_{H^\pm}$  in a two-doublet model  $m_{H^\pm} > 0.97 \tan\beta \text{ GeV}$  (90% CL).
- <sup>36</sup> COARASA 97 reanalyzed the constraint on the  $(m_{H^\pm}, \tan\beta)$  plane derived from the inclusive  $B \rightarrow \tau\nu_\tau X$  branching ratio in GROSSMAN 95B and BUSKULIC 95. They show that the constraint is quite sensitive to supersymmetric one-loop effects.
- <sup>37</sup> GUCHAIT 97 studies the constraints on  $m_{H^\pm}$  set by Tevatron data on  $\ell\tau$  final states in  $t\bar{t} \rightarrow (Wb)(Hb), W \rightarrow \ell\nu, H \rightarrow \tau\nu_\tau$ . See Fig. 2 for the excluded region.
- <sup>38</sup> MANGANO 97 reconsiders the limit in ACCIARRI 97F including the effect of the potentially large  $B_c \rightarrow \tau\nu_\tau$  background to  $B_u \rightarrow \tau\nu_\tau$  decays. Stronger limits are obtained.
- <sup>39</sup> STAHL 97 fit  $\tau$  lifetime, leptonic branching ratios, and the Michel parameters and derive limit  $m_{H^\pm} > 1.5 \tan\beta \text{ GeV}$  (90% CL) for a two-doublet model. See also STAHL 94.
- <sup>40</sup> ALAM 95 measure the inclusive  $b \rightarrow s\gamma$  branching ratio at  $T(45)$  and give  $B(b \rightarrow s\gamma) < 4.2 \times 10^{-4}$  (95% CL), which translates to the limit  $m_{H^\pm} > [244 + 63/(\tan\beta)^{1.3}] \text{ GeV}$  in the Type II two-doublet model. Light supersymmetric particles can invalidate this bound.
- <sup>41</sup> BUSKULIC 95 give a limit  $m_{H^\pm} > 1.9 \tan\beta \text{ GeV}$  (90% CL) for Type-II models from  $b \rightarrow \tau\nu_\tau X$  branching ratio, as proposed in GROSSMAN 94.

### MASS LIMITS for $H^{\pm\pm}$ (doubly-charged Higgs boson)

This section covers searches for a doubly-charged Higgs boson with couplings to lepton pairs. Its weak isospin  $T_3$  is thus restricted to two possibilities depending on lepton chiralities:  $T_3(H^{\pm\pm}) = \pm 1$ , with the coupling  $g_{\ell\ell}$  to  $\ell_L^-\ell_L^-$  and  $\ell_R^+\ell_R^+$  ("left-handed") and  $T_3(H^{\pm\pm}) = 0$ , with the coupling to  $\ell_R^-\ell_R^-$  and  $\ell_L^+\ell_L^+$  ("right-handed"). These Higgs bosons appear in some left-right symmetric models based on the gauge group  $SU(2)_L \times SU(2)_R \times U(1)$ , the type-II seesaw model, and the Zee-Babu model. The two cases are listed separately in the following. Unless noted, one of the lepton flavor combinations is assumed to be dominant in the decay.

### LIMITS for $H^{\pm\pm}$ with $T_3 = \pm 1$

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>551	95	1 AAD	15AG ATLS	$ee$
>468	95	1 AAD	15AG ATLS	$e\mu$
>516	95	1 AAD	15AG ATLS	$\mu\mu$
>400	95	2 AAD	15AP ATLS	$e\tau$
>400	95	2 AAD	15AP ATLS	$\mu\tau$
>169	95	3 CHATRCHYAN12AU	CMS	$\tau\tau$
>300	95	3 CHATRCHYAN12AU	CMS	$\mu\tau$
>293	95	3 CHATRCHYAN12AU	CMS	$e\tau$
>395	95	3 CHATRCHYAN12AU	CMS	$\mu\mu$
>391	95	3 CHATRCHYAN12AU	CMS	$e\mu$
>382	95	3 CHATRCHYAN12AU	CMS	$ee$
> 98.1	95	4 ABDALLAH 03	DLPH	$\tau\tau$
> 99.0	95	5 ABBIENDI 02c	OPAL	$\tau\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		6 KANEMURA 15	RVUE	$W^{(*)\pm}W^{(*)\pm}$
		7 KHACHATRYAN15D	CMS	$W^\pm W^\pm$
		8 KANEMURA 14	RVUE	$W^{(*)\pm}W^{(*)\pm}$
>330	95	9 AAD	13Y ATLS	$\mu\mu$
>237	95	9 AAD	13Y ATLS	$\mu\tau$
>355	95	10 AAD	12AY ATLS	$\mu\mu$
>398	95	11 AAD	12CQ ATLS	$\mu\mu$
>375	95	11 AAD	12CQ ATLS	$e\mu$
>409	95	11 AAD	12CQ ATLS	$ee$
>128	95	12 ABAZOV 12A	D0	$\tau\tau$
>144	95	12 ABAZOV 12A	D0	$\mu\tau$
>245	95	13 AALTONEN 11AF	CDF	$\mu\mu$
>210	95	13 AALTONEN 11AF	CDF	$e\mu$
>225	95	13 AALTONEN 11AF	CDF	$ee$
>114	95	14 AALTONEN 08AA	CDF	$e\tau$
>112	95	14 AALTONEN 08AA	CDF	$\mu\tau$
>168	95	15 ABAZOV 08V	D0	$\mu\mu$
		16 AKTAS 06A	H1	single $H^{\pm\pm}$
>133	95	17 ACOSTA 05L	CDF	stable
>118.4	95	18 ABAZOV 04E	D0	$\mu\mu$
		19 ABBIENDI 03Q	OPAL	$E_{\text{cm}} \leq 209 \text{ GeV}$ , single $H^{\pm\pm}$ muonium conversion
		20 GORDEEV 97	SPEC	
		21 ASAKA 95	THEO	
> 45.6	95	22 ACTON 92M	OPAL	
> 30.4	95	23 ACTON 92M	OPAL	
none 6.5–36.6	95	24 SWARTZ 90	MRK2	

- <sup>1</sup> AAD 15AG search for  $H^{++}H^{--}$  production in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8 \text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state. See their Fig. 5 for limits for arbitrary branching ratios.
- <sup>2</sup> AAD 15AP search for  $H^{++}H^{--}$  production in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8 \text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state.
- <sup>3</sup> CHATRCHYAN 12AU search for  $H^{++}H^{--}$  production with  $4.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state. See their Table 6 for limits including associated  $H^{++}H^{--}$  production or assuming different scenarios.
- <sup>4</sup> ABDALLAH 03 search for  $H^{++}H^{--}$  pair production either followed by  $H^{++} \rightarrow \tau^+\tau^+$ , or decaying outside the detector.
- <sup>5</sup> ABBIENDI 02c searches for pair production of  $H^{++}H^{--}$ , with  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$  ( $\ell, \ell' = e, \mu, \tau$ ). The limit holds for  $\ell=\ell'=\tau$ , and becomes stronger for other combinations of leptonic final states. To ensure the decay within the detector, the limit only applies for  $g(H\ell\ell) \gtrsim 10^{-7}$ .
- <sup>6</sup> KANEMURA 15 examine the case where  $H^{++}$  decays preferentially to  $W^{(*)}W^{(*)}$  and estimate that a lower mass limit of  $\sim 84 \text{ GeV}$  can be derived from the same-sign dilepton data of AAD 15AG if  $H^{++}$  decays with 100% branching ratio to  $W^{(*)}W^{(*)}$ .
- <sup>7</sup> KHACHATRYAN 15D search for  $H^{\pm\pm}$  production by vector boson fusion followed by the decay  $H^{\pm\pm} \rightarrow W^\pm W^\pm$  in  $19.4 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 8 \text{ TeV}$ . See their Fig. 4 for limits on cross section times branching ratio for  $m_{H^{++}}$  between 160 and 800 GeV.
- <sup>8</sup> KANEMURA 14 examine the case where  $H^{++}$  decays preferentially to  $W^{(*)}W^{(*)}$  and estimate that a lower mass limit of  $\sim 60 \text{ GeV}$  can be derived from the same-sign dilepton data of AAD 12CY.
- <sup>9</sup> AAD 13Y search for  $H^{++}H^{--}$  production in a generic search of events with three charged leptons in  $4.6 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state.
- <sup>10</sup> AAD 12AY search for  $H^{++}H^{--}$  production with  $1.6 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state.
- <sup>11</sup> AAD 12CQ search for  $H^{++}H^{--}$  production with  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $E_{\text{cm}} = 7 \text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state. See their Table 1 for limits assuming smaller branching ratios.

# Gauge & Higgs Boson Particle Listings

## Charged Higgs Bosons ( $H^\pm$ and $H^{\pm\pm}$ ), Searches for

- <sup>12</sup> ABAZOV 12A search for  $H^{++}H^{--}$  production in  $7.0\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ .
- <sup>13</sup> AALTONEN 11AF search for  $H^{++}H^{--}$  production in  $6.1\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ .
- <sup>14</sup> AALTONEN 08AA search for  $H^{++}H^{--}$  production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state.
- <sup>15</sup> ABAZOV 08v search for  $H^{++}H^{--}$  production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ . The limit is for  $B(H \rightarrow \mu\mu) = 1$ . The limit is updated in ABAZOV 12A.
- <sup>16</sup> AKTAS 06A search for single  $H^{\pm\pm}$  production in  $ep$  collisions at HERA. Assuming that  $H^{++}$  only couples to  $e^+\mu^+$  with  $g_{e\mu} = 0.3$  (electromagnetic strength), a limit  $m_{H^{++}} > 141\text{ GeV}$  (95% CL) is derived. For the case where  $H^{++}$  couples to  $e\tau$  only the limit is 112 GeV.
- <sup>17</sup> ACOSTA 05L search for  $H^{++}H^{--}$  pair production in  $p\bar{p}$  collisions. The limit is valid for  $g_{\ell\ell'} < 10^{-8}$  so that the Higgs decays outside the detector.
- <sup>18</sup> ABAZOV 04E search for  $H^{++}H^{--}$  pair production in  $H^{\pm\pm} \rightarrow \mu^\pm\mu^\pm$ . The limit is valid for  $g_{\mu\mu} \gtrsim 10^{-7}$ .
- <sup>19</sup> ABBIENDI 03Q searches for single  $H^{\pm\pm}$  via direct production in  $e^+e^- \rightarrow e^\mp e^\mp H^{\pm\pm}$ , and via  $t$ -channel exchange in  $e^+e^- \rightarrow e^+e^-$ . In the direct case, and assuming  $B(H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm) = 1$ , a 95% CL limit on  $h_{ee} < 0.071$  is set for  $m_{H^{\pm\pm}} < 160\text{ GeV}$  (see Fig. 6). In the second case, indirect limits on  $h_{ee}$  are set for  $m_{H^{\pm\pm}} < 2\text{ TeV}$  (see Fig. 8).
- <sup>20</sup> GORDEEV 97 search for muonium-antimuonium conversion and find  $G_{M\bar{M}}/G_F < 0.14$  (90% CL), where  $G_{M\bar{M}}$  is the lepton-flavor violating effective four-fermion coupling. This limit may be converted to  $m_{H^{++}} > 210\text{ GeV}$  if the Yukawa couplings of  $H^{++}$  to  $ee$  and  $\mu\mu$  are as large as the weak gauge coupling. For similar limits on muonium-antimuonium conversion, see the muon Particle Listings.
- <sup>21</sup> ASAKA 95 point out that  $H^{++}$  decays dominantly to four fermions in a large region of parameter space where the limit of ACTON 92M from the search of dilepton modes does not apply.
- <sup>22</sup> ACTON 92M limit assumes  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$  or  $H^{\pm\pm}$  does not decay in the detector. Thus the region  $g_{\ell\ell} \approx 10^{-7}$  is not excluded.
- <sup>23</sup> ACTON 92M from  $\Delta\Gamma_Z < 40\text{ MeV}$ .
- <sup>24</sup> SWARTZ 90 assume  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$  (any flavor). The limits are valid for the Higgs-lepton coupling  $g(H\ell\ell) \gtrsim 7.4 \times 10^{-7}/[m_H/\text{GeV}]^{1/2}$ . The limits improve somewhat for  $ee$  and  $\mu\mu$  decay modes.

### LIMITS FOR $H^{\pm\pm}$ with $T_3 = 0$

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>374	95	1 AAD	15AG ATLS	$ee$
>402	95	1 AAD	15AG ATLS	$e\mu$
>438	95	1 AAD	15AG ATLS	$\mu\mu$
>290	95	2 AAD	15AP ATLS	$e\tau$
>290	95	2 AAD	15AP ATLS	$\mu\tau$
> 97.3	95	3 ABDALLAH	03 DLPH	$\tau\tau$
> 97.3	95	4 ACHARD	03F L3	$\tau\tau$
> 98.5	95	5 ABBIENDI	02C OPAL	$\tau\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>251	95	6 AAD	12AY ATLS	$\mu\mu$
>306	95	7 AAD	12CQ ATLS	$\mu\mu$
>310	95	7 AAD	12CQ ATLS	$e\mu$
>322	95	7 AAD	12CQ ATLS	$ee$
>113	95	8 ABAZOV	12A D0	$\mu\tau$
>205	95	9 AALTONEN	11AF CDF	$\mu\mu$
>190	95	9 AALTONEN	11AF CDF	$e\mu$
>205	95	9 AALTONEN	11AF CDF	$ee$
>145	95	10 ABAZOV	08v D0	$\mu\mu$
>109	95	11 AKTAS	06A H1	single $H^{\pm\pm}$
> 98.2	95	12 ACOSTA	05L CDF	stable
		13 ABAZOV	04E D0	$\mu\mu$
		14 ABBIENDI	03Q OPAL	$E_{\text{cm}} \leq 209\text{ GeV}$ , single $H^{\pm\pm}$
		15 GORDEEV	97 SPEC	muonium conversion
> 45.6	95	16 ACTON	92M OPAL	
> 25.5	95	17 ACTON	92M OPAL	
none 7.3–34.3	95	18 SWARTZ	90 MRK2	

- <sup>1</sup> AAD 15AG search for  $H^{++}H^{--}$  production in  $20.3\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state. See their Fig. 5 for limits for arbitrary branching ratios.
- <sup>2</sup> AAD 15AP search for  $H^{++}H^{--}$  production in  $20.3\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state.
- <sup>3</sup> ABDALLAH 03 search for  $H^{++}H^{--}$  pair production either followed by  $H^{++} \rightarrow \tau^+\tau^+$ , or decaying outside the detector.
- <sup>4</sup> ACHARD 03F search for  $e^+e^- \rightarrow H^{++}H^{--}$  with  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$ . The limit holds for  $\ell = \ell' = \tau$ , and slightly different limits apply for other flavor combinations. The limit is valid for  $g_{\ell\ell'} \gtrsim 10^{-7}$ .
- <sup>5</sup> ABBIENDI 02c searches for pair production of  $H^{++}H^{--}$ , with  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$  ( $\ell, \ell' = e, \mu, \tau$ ). the limit holds for  $\ell = \ell' = \tau$ , and becomes stronger for other combinations of leptonic final states. To ensure the decay within the detector, the limit only applies for  $g(H\ell\ell) \gtrsim 10^{-7}$ .
- <sup>6</sup> AAD 12AY search for  $H^{++}H^{--}$  production with  $1.6\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state.
- <sup>7</sup> AAD 12CQ search for  $H^{++}H^{--}$  production with  $4.7\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ . The limit assumes 100% branching ratio to the specified final state. See their Table 1 for limits assuming smaller branching ratios.
- <sup>8</sup> ABAZOV 12A search for  $H^{++}H^{--}$  production in  $7.0\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ .

- <sup>9</sup> AALTONEN 11AF search for  $H^{++}H^{--}$  production in  $6.1\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ .
- <sup>10</sup> ABAZOV 08v search for  $H^{++}H^{--}$  production in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96\text{ TeV}$ . The limit is for  $B(H \rightarrow \mu\mu) = 1$ . The limit is updated in ABAZOV 12A.
- <sup>11</sup> AKTAS 06A search for single  $H^{\pm\pm}$  production in  $ep$  collisions at HERA. Assuming that  $H^{++}$  only couples to  $e^+\mu^+$  with  $g_{e\mu} = 0.3$  (electromagnetic strength), a limit  $m_{H^{++}} > 141\text{ GeV}$  (95% CL) is derived. For the case where  $H^{++}$  couples to  $e\tau$  only the limit is 112 GeV.
- <sup>12</sup> ACOSTA 05L search for  $H^{++}H^{--}$  pair production in  $p\bar{p}$  collisions. The limit is valid for  $g_{\ell\ell'} < 10^{-8}$  so that the Higgs decays outside the detector.
- <sup>13</sup> ABAZOV 04E search for  $H^{++}H^{--}$  pair production in  $H^{\pm\pm} \rightarrow \mu^\pm\mu^\pm$ . The limit is valid for  $g_{\mu\mu} \gtrsim 10^{-7}$ .
- <sup>14</sup> ABBIENDI 03Q searches for single  $H^{\pm\pm}$  via direct production in  $e^+e^- \rightarrow e^\mp e^\mp H^{\pm\pm}$ , and via  $t$ -channel exchange in  $e^+e^- \rightarrow e^+e^-$ . In the direct case, and assuming  $B(H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm) = 1$ , a 95% CL limit on  $h_{ee} < 0.071$  is set for  $m_{H^{\pm\pm}} < 160\text{ GeV}$  (see Fig. 6). In the second case, indirect limits on  $h_{ee}$  are set for  $m_{H^{\pm\pm}} < 2\text{ TeV}$  (see Fig. 8).
- <sup>15</sup> GORDEEV 97 search for muonium-antimuonium conversion and find  $G_{M\bar{M}}/G_F < 0.14$  (90% CL), where  $G_{M\bar{M}}$  is the lepton-flavor violating effective four-fermion coupling. This limit may be converted to  $m_{H^{++}} > 210\text{ GeV}$  if the Yukawa couplings of  $H^{++}$  to  $ee$  and  $\mu\mu$  are as large as the weak gauge coupling. For similar limits on muonium-antimuonium conversion, see the muon Particle Listings.
- <sup>16</sup> ACTON 92M limit assumes  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$  or  $H^{\pm\pm}$  does not decay in the detector. Thus the region  $g_{\ell\ell} \approx 10^{-7}$  is not excluded.
- <sup>17</sup> ACTON 92M from  $\Delta\Gamma_Z < 40\text{ MeV}$ .
- <sup>18</sup> SWARTZ 90 assume  $H^{\pm\pm} \rightarrow \ell^\pm\ell^\pm$  (any flavor). The limits are valid for the Higgs-lepton coupling  $g(H\ell\ell) \gtrsim 7.4 \times 10^{-7}/[m_H/\text{GeV}]^{1/2}$ . The limits improve somewhat for  $ee$  and  $\mu\mu$  decay modes.

### $H^\pm$ and $H^{\pm\pm}$ REFERENCES

AAD	15AF	JHEP 1503 088	G. Aad et al.	(ATLAS Collab.)
AAD	15AG	JHEP 1503 041	G. Aad et al.	(ATLAS Collab.)
AAD	15AP	JHEP 1508 138	G. Aad et al.	(ATLAS Collab.)
AAD	15M	PRL 114 231801	G. Aad et al.	(ATLAS Collab.)
KANEMURA	15	PTEP 2015 051B02	S. Kanemura et al.	
KHACHATRYAN	15AX	JHEP 1511 018	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15BF	JHEP 1512 178	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15D	PRL 114 051801	V. Khachatryan et al.	(CMS Collab.)
AAD	14M	PR D89 032002	G. Aad et al.	(ATLAS Collab.)
AALTONEN	14A	PR D89 091101	T. Aaltonen et al.	(CDF Collab.)
KANEMURA	14	PR D90 115018	S. Kanemura et al.	
AAD	13AC	EPJ C73 2465	G. Aad et al.	(ATLAS Collab.)
AAD	13V	JHEP 1303 076	G. Aad et al.	(ATLAS Collab.)
AAD	13Y	PR D87 052002	G. Aad et al.	(ATLAS Collab.)
LEP	13	EPJ C73 2463	LEP Collabs	(ALEPH, DELPHI, L3, OPAL, LEP)
AAD	12AY	PR D85 032004	G. Aad et al.	(ATLAS Collab.)
AAD	12BH	JHEP 1206 039	G. Aad et al.	(ATLAS Collab.)
AAD	12CQ	EPJ C72 2244	G. Aad et al.	(ATLAS Collab.)
AAD	12CY	JHEP 1212 007	G. Aad et al.	(ATLAS Collab.)
ABAZOV	12A	PRL 108 021801	V.M. Abazov et al.	(D0 Collab.)
ABBIENDI	12	EPJ C72 2076	G. Abbiendi et al.	(OPAL Collab.)
CHATRCHYAN	12AA	JHEP 1207 143	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12AU	EPJ C72 2189	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN	11AF	PRL 107 181801	T. Aaltonen et al.	(CDF Collab.)
DESCHAMPS	11P	PRL 107 031801	T. Aaltonen et al.	(CDF Collab.)
ACTON	10	PR D82 073012	O. Deschamps et al.	(CLER, ORSAY, LAPP)
AALTONEN	09AJ	PRL 103 101803	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	09AC	PR D80 051107	V.M. Abazov et al.	(D0 Collab.)
ABAZOV	09AG	PR D80 071102	V.M. Abazov et al.	(D0 Collab.)
ABAZOV	09AI	PL B682 278	V.M. Abazov et al.	(D0 Collab.)
ABAZOV	09P	PRL 102 191802	V.M. Abazov et al.	(D0 Collab.)
AALTONEN	08AA	PRL 101 121801	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	08V	PRL 101 071803	V.M. Abazov et al.	(D0 Collab.)
ABULENCIA	06E	PRL 96 042003	A. Abulencia et al.	(CDF Collab.)
AKTAS	06A	PL B638 432	A. Aktas et al.	(H1 Collab.)
ACOSTA	05L	PRL 95 071801	D. Acosta et al.	(CDF Collab.)
ABAZOV	04E	PRL 93 141801	V.M. Abazov et al.	(D0 Collab.)
ABBIENDI	04	EPJ C32 453	G. Abbiendi et al.	(OPAL Collab.)
ABDALLAH	04I	EPJ C34 399	J. Abdallah et al.	(DELPHI Collab.)
ABBIENDI	03	PL B551 35	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI	03Q	PL B577 93	G. Abbiendi et al.	(OPAL Collab.)
ABDALLAH	03	PL B552 127	J. Abdallah et al.	(DELPHI Collab.)
ACHARD	03E	PL B575 208	P. Achard et al.	(L3 Collab.)
ACHARD	03F	PL B576 18	P. Achard et al.	(L3 Collab.)
ABAZOV	02B	PRL 88 151803	V.M. Abazov et al.	(D0 Collab.)
ABBIENDI	02C	PL B526 221	G. Abbiendi et al.	(OPAL Collab.)
BORZUMATI	02	PL B549 170	F.M. Borzumati, A. Djouadi	
HEISTER	02P	PL B543 1	A. Heister et al.	(ALEPH Collab.)
ABBIENDI	01Q	PL B520 1	G. Abbiendi et al.	(OPAL Collab.)
BARATE	01E	EPJ C19 213	R. Barate et al.	(ALEPH Collab.)
GAMBINO	01	NP B611 338	P. Gambino, M. Misiak	
AFFOLDER	00I	PR D62 012004	T. Affolder et al.	(CDF Collab.)
PDG	00	EPJ C15 1	D.E. Groom et al.	(PDG Collab.)
ABBIENDI	99E	EPJ C7 407	G. Abbiendi et al.	(OPAL Collab.)
ABBOTT	99E	PRL 82 4975	B. Abbott et al.	(D0 Collab.)
ACKERSTAFF	99D	EPJ C8 3	K. Ackerstaff et al.	(OPAL Collab.)
ABE	97L	PRL 79 357	F. ABE et al.	(CDF Collab.)
ACCIARRI	97F	PL B396 327	M. Acciarri et al.	(L3 Collab.)
AMMAR	97B	PRL 78 4686	R. Ammar et al.	(CLEO Collab.)
COARASA	97	PL B406 337	J.A. Coarasa, R.A. Jimenez, J. Sola	
GORDEEV	97	PAN 60 1164	V.A. Gordeev et al.	(PNPI)
		Translated from YAF 60 1291.		
GUCHAIT	97	PR D55 7263	M. Guhait, D.P. Roy	(TATA)
MANGANO	97	PR D410 299	M. Mangano, S. Slabospitsky	
STAHL	97	ZPHY C74 73	A. Stahl, H. Voss	(BONN)
PDG	96	PL B54 1	R.M. Barnett et al.	(PDG Collab.)
ALAM	95	PRL 74 2885	M.S. Alam et al.	(CLEO Collab.)
ASAKA	95	PL B345 36	T. Asaka, K.I. Hikasa	(TOHOKU)
BUSKULIC	95	PL B343 444	D. Buskulic et al.	(ALEPH Collab.)
GROSSMAN	95B	PL B357 630	Y. Grossman, H. Haber, Y. Nir	
GROSSMAN	94	PL B332 373	Y. Grossman, Z. Ligeti	
STAHL	94	PL B324 121	A. Stahl	(BONN)
ACTON	92M	PL B295 347	P.D. Acton et al.	(OPAL Collab.)
SWARTZ	90	PRL 64 2877	M.L. Swartz et al.	(Mark II Collab.)



## New Heavy Bosons ( $W'$ , $Z'$ , leptoquarks, etc.), Searches for

We list here various limits on charged and neutral heavy vector bosons (other than  $W$ 's and  $Z$ 's), heavy scalar bosons (other than Higgs bosons), vector or scalar leptoquarks, and axigluons. The latest unpublished results are described in “ $W'$  Searches” and “ $Z'$  Searches” reviews. For recent searches on scalar bosons which could be identified as Higgs bosons, see the listings in the Higgs boson section.

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 – Limits for  $Z_{LR}^{SM}$   
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### $W'$ -BOSON SEARCHES

Revised March 2016 by M.-C. Chen (UC Irvine), B.A. Dobrescu (Fermilab) and S. Willocq (U Massachusetts).

The  $W'$  boson is a massive hypothetical particle of spin 1 and electric charge  $\pm 1$ , which is a color singlet and is predicted in various extensions of the Standard Model (SM).

**$W'$  couplings to quarks and leptons.** The Lagrangian terms describing couplings of a  $W'^+$  boson to fermions are given by

$$\frac{W'^+_\mu}{\sqrt{2}} \left[ \bar{u}_i (C_{qij}^R P_R + C_{qij}^L P_L) \gamma^\mu d_j + \bar{\nu}_i (C_{\ell ij}^R P_R + C_{\ell ij}^L P_L) \gamma^\mu e_j \right]. \quad (1)$$

Here  $u, d, \nu$  and  $e$  are the SM fermions in the mass eigenstate basis,  $i, j = 1, 2, 3$  label the fermion generation, and  $P_{R,L} = (1 \pm \gamma_5)/2$ . The coefficients  $C_{qij}^R$ ,  $C_{qij}^L$ ,  $C_{\ell ij}^R$ , and  $C_{\ell ij}^L$  are complex dimensionless parameters. If  $C_{\ell ij}^R \neq 0$ , then the  $i$ th generation includes a right-handed neutrino. Using this notation, the SM  $W$  couplings are  $C_q^L = g_{VCKM}$ ,  $C_\ell^L = g \approx 0.63$  and  $C_q^R = C_\ell^R = 0$ .

Unitarity considerations imply that the  $W'$  boson is associated with a spontaneously-broken gauge symmetry. This is true even when it is a composite particle (*e.g.*,  $\rho^\pm$ -like bound states [1]) if its mass is much smaller than the compositeness scale, or a Kaluza-Klein mode in theories where the  $W$  boson propagates in extra dimensions [2]. The simplest extension of the electroweak gauge group that includes a  $W'$  boson is  $SU(2)_1 \times SU(2)_2 \times U(1)$ , but larger groups are encountered in some theories. A generic property of these gauge theories is

that they also include a  $Z'$  boson [3]; whether the  $W'$  boson can be discovered first depends on theoretical and experimental details.

A tree-level mass mixing may be induced between the electrically-charged gauge bosons. Upon diagonalization of their mass matrix, the  $W - Z$  mass ratio and the couplings of the observed  $W$  boson are shifted from the SM values. Their measurements imply that the mixing angle between the gauge eigenstates,  $\theta_+$ , must be smaller than about  $10^{-2}$ . In certain theories the mixing is negligible (*e.g.* due to a new parity [4]), even when the  $W'$  mass is near the electroweak scale.

The  $W'$  coupling to  $WZ$  is fixed by Lorentz and gauge invariances, and to leading order in  $\theta_+$  is given by [5]

$$\frac{g \theta_+ i}{\cos \theta_W} [W'^+_\mu (W^-_\nu Z^{\nu\mu} + Z_\nu W^{-\mu\nu}) + Z^\nu W^{-\mu} W'^+_{\nu\mu}] + \text{H.c.}, \quad (2)$$

where  $W^{\mu\nu} \equiv \partial^\mu W^\nu - \partial^\nu W^\mu$ , etc. The  $\theta_W$  dependence shown here corrects the one given in [6], which has been referred to as the Extended Gauge Model by the experimental collaborations. The  $W'$  coupling to  $Wh^0$ , where  $h^0$  is the SM Higgs boson, is

$$-\xi_h g_{W'} M_W W'^+_\mu W^{\mu-} h^0 + \text{H.c.}, \quad (3)$$

where  $g_{W'}$  is the gauge coupling of the  $W'$  boson, and the coefficient  $\xi_h$  satisfies  $\xi_h \leq 1$  in simple Higgs sectors [5].

In models based on the “left-right symmetric” gauge group [7],  $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$ , the SM fermions that couple to the  $W$  boson transform as doublets under  $SU(2)_L$  while the other fermions transform as doublets under  $SU(2)_R$ . Consequently, the  $W'$  boson couples primarily to right-handed fermions; its coupling to left-handed fermions arises due to the  $\theta_+$  mixing, so that  $C_q^L$  is proportional to the CKM matrix and its elements are much smaller than the diagonal elements of  $C_q^R$ . Generically,  $C_q^R$  does not need to be proportional to  $V_{CKM}$ .

There are many other models based on the  $SU(2)_1 \times SU(2)_2 \times U(1)$  gauge symmetry. In the “alternate left-right” model [8], all the couplings shown in Eq. (1) vanish, but there are some new fermions such that the  $W'$  boson couples to pairs involving a SM fermion and a new fermion. In the “unified SM” [9], the left-handed quarks are doublets under one  $SU(2)$ , and the left-handed leptons are doublets under a different  $SU(2)$ , leading to a mostly leptophobic  $W'$  boson:  $C_{\ell ij}^L \ll C_{qij}^L$  and  $C_{\ell ij}^R = C_{qij}^R = 0$ . Fermions of different generations may also transform as doublets under different  $SU(2)$  gauge groups [10]. In particular, the couplings to third generation quarks may be enhanced [11].

It is also possible that the  $W'$  couplings to SM fermions are highly suppressed. For example, if the quarks and leptons are singlets under one  $SU(2)$  [12], then the couplings are proportional to the tiny mixing angle  $\theta_+$ . Similar suppressions may arise if some vectorlike fermions mix with the SM fermions [13].

Gauge groups that embed the electroweak symmetry, such as  $SU(3)_W \times U(1)$  or  $SU(4)_W \times U(1)$ , also include one or more  $W'$  bosons [14].

## Gauge & Higgs Boson Particle Listings

### New Heavy Bosons

**Collider searches.** At LEP-II,  $W'$  bosons could have been produced in pairs via their photon and  $Z$  couplings. The production cross section is large enough to rule out  $M_{W'} < \sqrt{s}/2 \approx 105$  GeV for most patterns of decay modes.

At hadron colliders,  $W'$  bosons can be detected through resonant pair production of fermions or electroweak bosons. Assuming that the  $W'$  width is much smaller than its mass, the contribution of the  $s$ -channel  $W'$  boson exchange to the total rate for  $pp \rightarrow f\bar{f}'X$ , where  $f$  and  $f'$  are fermions with an  $f\bar{f}'$  electric charge of  $\pm 1$ , and  $X$  is any final state, may be approximated by the branching fraction  $B(W' \rightarrow f\bar{f}')$  times the production cross section

$$\sigma(pp \rightarrow W'X) \simeq \frac{\pi}{48s} \sum_{i,j} \left[ (C_{qij}^L)^2 + (C_{qij}^R)^2 \right] w_{ij}(M_{W'}^2/s, M_{W'}) . \quad (4)$$

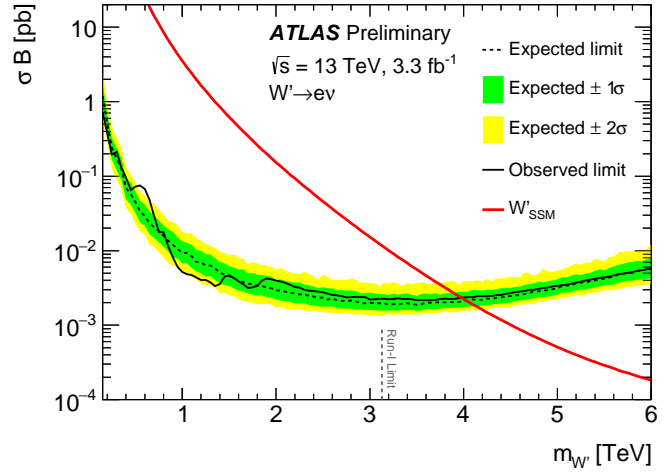
The functions  $w_{ij}$  include the information about proton structure, and are given to leading order in  $\alpha_s$  by

$$w_{ij}(z, \mu) = \int_z^1 \frac{dx}{x} \left[ u_i(x, \mu) \bar{d}_j\left(\frac{z}{x}, \mu\right) + \bar{u}_i(x, \mu) d_j\left(\frac{z}{x}, \mu\right) \right], \quad (5)$$

where  $u_i(x, \mu)$  and  $d_i(x, \mu)$  are the parton distributions inside the proton, at the factorization scale  $\mu$  and parton momentum fraction  $x$ , for the up- and down-type quark of the  $i$ th generation, respectively. QCD corrections to  $W'$  production are sizable (they also include quark-gluon initial states), but preserve the above factorization of couplings at next-to-leading order [15].

The most commonly studied  $W'$  signal consists of a high-momentum electron or muon and large missing transverse momentum, with the transverse mass distribution forming a Jacobian peak with its endpoint at  $M_{W'}$  (see Fig. 1e of [16]). Given that the branching fractions for  $W' \rightarrow e\nu$  and  $W' \rightarrow \mu\nu$  could be very different, these channels should be analyzed separately. Searches in these channels often implicitly assume that the left-handed couplings vanish (no interference between  $W$  and  $W'$ ), and that the right-handed neutrino is light compared to the  $W'$  boson and escapes the detector. These assumptions correspond to the following choice of parameters:  $C_q^R = gV_{\text{CKM}}$ ,  $C_q^L = g$ ,  $C_\ell^L = C_\ell^R = 0$ , which define a model that is essentially equivalent to the Sequential SM used in many searches. However, if a  $W'$  boson were discovered and the final state fermions have left-handed helicity, then the effects of  $W - W'$  interference could be observed [17], providing useful information about the  $W'$  couplings.

In the  $e\nu$  channel, the ATLAS and CMS Collaborations set limits on the  $W'$  production cross section times branching fraction (and thus indirectly on the  $W'$  couplings) when  $M_{W'}$  is in the 0.2 – 6 TeV range, based on 20 fb $^{-1}$  of LHC data at  $\sqrt{s} = 8$  TeV [16,18] and 2–3 fb $^{-1}$  at  $\sqrt{s} = 13$  TeV [19,20], as shown in Fig. 1. ATLAS sets the strongest mass lower limit  $M_{W'} > 4.0$  TeV in the Sequential SM (all limits in this mini-review are at the 95% CL). The coupling limits are much weaker for  $M_{W'} < 200$  GeV, a range last explored with the Tevatron at  $\sqrt{s} = 1.8$  TeV [21].



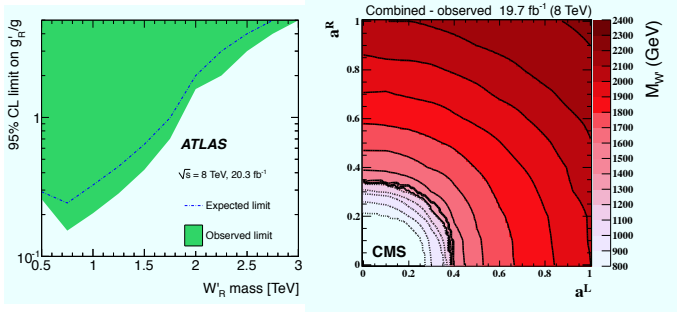
**Figure 1:** Upper limit on  $\sigma(pp \rightarrow W'X) B(W' \rightarrow e\nu)$  from ATLAS [20], at 95% CL. The red line shows the theoretical prediction in the Sequential SM.

In the  $\mu\nu$  channel, ATLAS and CMS set rate limits for  $M_{W'}$  in the 0.2 – 6 TeV range from the same analyses as mentioned above, with the strongest lower mass limit of 4.0 TeV set by CMS [19] using the  $\sqrt{s} = 13$  TeV data. When combined with the  $e\nu$  channel, the upper limit on the  $\sqrt{s} = 13$  TeV cross section times branching fraction to  $\ell\nu$  varies between 1 and 2 fb for  $M_{W'}$  between 1 and 5 TeV [19]. Only weak limits on  $W' \rightarrow \mu\nu$  exist for  $M_{W'} < 200$  GeV [22]. Note that masses of the order of the electroweak scale are interesting from a theory point of view, while lepton universality does not necessarily apply to a  $W'$  boson.

A dedicated search for  $W' \rightarrow \tau\nu$  has been performed by the CMS Collaboration at 8 TeV [23]. Limits are set on  $\sigma \cdot B$  for  $M_{W'}$  between 0.3 and 4.0 TeV. A lower mass limit of 2.7 TeV is set in the Sequential SM.

The  $W'$  decay into a lepton and a right-handed neutrino,  $\nu_R$ , may also be followed by the  $\nu_R$  decay through a virtual  $W'$  boson into a lepton and two quark jets. The ATLAS [24] and CMS [25] searches in the  $eejj$  and  $\mu\mu jj$  channels have set limits on the cross section times branching fraction as a function of the  $\nu_R$  mass or of  $M_{W'}$ . These searches are typically performed with same-charge lepton pairs that provide strong background reduction and are motivated by models with a left-right symmetry. However, it is also interesting to search in final states with opposite-charge lepton pairs, as done in the CMS analysis.

The  $t\bar{b}$  channel is particularly important because a  $W'$  boson that couples only to right-handed fermions cannot decay to leptons when the right-handed neutrinos are heavier than the  $W'$  boson (additional motivations are provided by a  $W'$  boson with enhanced couplings to the third generation [11], and by a leptophobic  $W'$  boson). The usual signature consists of a leptonically-decaying  $W$  boson and two  $b$ -jets. Recent studies have also incorporated the fully hadronic decay channel for  $M_{W'} \gg m_t$  with the use of jet substructure techniques to tag



**Figure 2:** Upper limits on  $W'$  couplings (at 95% CL) using the  $t\bar{b}$  and  $\bar{t}b$  final states, assuming that the diagonal couplings are generation independent. Left panel: ATLAS [26] limit on  $C_{q11}^R/g$ . Right panel: CMS [27] limit on  $M_{W'}$  as contours in the  $C_{q11}^R/g - C_{q11}^L/g$  plane.

highly boosted top-jets. Upper limits on the  $W'$  couplings to right- and left-handed quarks normalized to the SM  $W$  couplings have been set by ATLAS [26] and CMS [27] at  $\sqrt{s} = 8$  TeV, as shown in Fig. 2. Using about 2 fb<sup>-1</sup> of data at  $\sqrt{s} = 13$  TeV in the  $\ell + \text{jets}$  channel, CMS [28] sets an upper limit on the  $W'$  production cross section times branching fraction to the  $\ell\nu b\bar{b}$  final state decreasing from 1.6 pb at  $M_{W'} = 1$  TeV to 35 fb at  $M_{W'} = 3$  TeV. The limit  $M_{W'} > 2.38$  TeV obtained in the Sequential SM with a light  $\nu_R$  increases with the  $\nu_R$  mass. The best limits on the couplings to right-handed quarks for  $M_{W'}$  in the 300–600 GeV range have been set by CDF with 9.5 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV [29]. Finally, if  $W'$  couplings to left-handed quarks are large, then interference effects modify the SM  $s$ -channel single-top production [30].

Searches for dijet resonances may be used to set limits on  $W' \rightarrow q\bar{q}'$ . The best limits on  $W'$  couplings to quarks have been set by UA2 [31] in the 140 – 250 GeV mass range, by CDF [32] in the 250 – 500 GeV range and by CMS [33] in the 500 – 750 GeV range. ATLAS and CMS provide similar coverage in the  $\sim 0.75 - 7$  TeV range with data collected at  $\sqrt{s} = 8$  and 13 TeV [34] with the most stringent lower  $W'$  mass limit in the Sequential SM set to 2.6 TeV using 13 TeV data.

In some theories [4], the  $W'$  couplings to SM fermions are suppressed by discrete symmetries.  $W'$  production then occurs in pairs, through a photon or  $Z$  boson. The decay modes are model-dependent and often involve other new particles. The ensuing collider signals arise from cascade decays and typically include missing transverse momentum.

Searches for  $WZ$  resonances at the LHC have focused on the process  $pp \rightarrow W' \rightarrow WZ$  with the production mainly from  $u\bar{d} \rightarrow W'$  assuming SM-like couplings to quarks. ATLAS and CMS have set the strongest upper limits on the  $W'WZ$  coupling for  $M_{W'}$  in the 0.2 – 4 TeV range with a combination of fully leptonic, semi-leptonic and fully hadronic channels at both 8 and 13 TeV [35,36,37,38]. ATLAS has also combined the results from all channels at 8 TeV and obtains  $M_{W'} > 1.81$  TeV in the Sequential SM [39].

A fermiophobic  $W'$  boson that couples to  $WZ$  may be produced at hadron colliders in association with a  $Z$  boson, or

via  $WZ$  fusion. This would give rise to  $(WZ)Z$  and  $(WZ)jj$  final states, where the parentheses represent a resonance [40].

$W'$  bosons have also been searched for recently in final states with a  $W$  boson and a SM Higgs boson in the channels  $W \rightarrow \ell\nu$  and  $h^0 \rightarrow b\bar{b}$  or  $h^0 \rightarrow WW$  by ATLAS [41,42] and CMS [43] at  $\sqrt{s} = 8$  and 13 TeV. Cross section limits are set for  $W'$  masses in the range between 0.4 and 3.0 TeV. The strongest lower limit on the mass is set by the ATLAS 13 TeV analysis:  $M_{W'} > 1.49$  TeV in the context of the Heavy Vector Triplet weakly-coupled scenario A [44].

**Low-energy constraints.** The properties of  $W'$  bosons are also constrained by measurements of processes at energies much below  $M_{W'}$ . The bounds on  $W - W'$  mixing [45] are mostly due to the change in  $W$  properties compared to the SM. Limits on deviations in the  $ZWW$  couplings provide a leading constraint for fermiophobic  $W'$  bosons [13].

Constraints arising from low-energy effects of  $W'$  exchange are strongly model-dependent. If the  $W'$  couplings to quarks are not suppressed, then box diagrams involving a  $W$  and a  $W'$  boson contribute to neutral meson-mixing. In the case of  $W'$  couplings to right-handed quarks as in the left-right symmetric model, the limit from  $K_L - K_S$  mixing is severe:  $M_{W'} > 2.9$  TeV for  $C_q^L = C_q^R$  [46]. However, if no correlation between the  $W'$  and  $W$  couplings is assumed, then the limit on  $M_{W'}$  may be significantly relaxed [47].

$W'$  exchange also contributes at tree level to various low-energy processes. In particular, it would impact the measurement of the Fermi constant  $G_F$  in muon decay, which in turn would change the predictions of many other electroweak processes. A recent test of parity violation in polarized muon decay [48] has set limits of about 600 GeV on  $M_{W'}$ , assuming  $W'$  couplings to right-handed leptons as in left-right symmetric models and a light  $\nu_R$ . There are also  $W'$  contributions to the neutron electric dipole moment,  $\beta$  decays, and other processes [45].

If right-handed neutrinos have Majorana masses, then there are tree-level contributions to neutrinoless double-beta decay, and a limit on  $M_{W'}$  versus the  $\nu_R$  mass may be derived [49]. For  $\nu_R$  masses below a few GeV, the  $W'$  boson contributes to leptonic and semileptonic  $B$  meson decays, so that limits may be placed on various combinations of  $W'$  parameters [47]. For  $\nu_R$  masses below  $\sim 30$  MeV, the most stringent constraints on  $M_{W'}$  are due to the limits on  $\nu_R$  emission from supernovae.

## References

1. M. Bando, T. Kugo, and K. Yamawaki, Phys. Rept. **164**, 217 (1988).
2. H.C. Cheng *et al.*, Phys. Rev. D **64**, 065007 (2001).
3. See the Section on “ $Z'$ -boson searches” in this *Review*.
4. H.C. Cheng and I. Low, JHEP **0309**, 051 (2003).
5. B. A. Dobrescu and Z. Liu, JHEP **1510**, 118 (2015).
6. G. Altarelli, B. Mele and M. Ruiz-Altaba, Z. Phys. C **45**, 109 (1989) [Z. Phys. C **47**, 676 (1990)].

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

7. R.N. Mohapatra and J.C. Pati, Phys. Rev. D **11**, 566 (1975); G. Senjanovic and R.N. Mohapatra, Phys. Rev. D **12**, 1502 (1975).
8. K.S. Babu, X.G. He, and E. Ma, Phys. Rev. D **36**, 878 (1987).
9. H. Georgi, E.E. Jenkins, and E.H. Simmons, Nucl. Phys. B **331**, 541 (1990).
10. See, *e.g.*, X. Li and E. Ma, J. Phys. G **19**, 1265 (1993).
11. D.J. Muller and S. Nandi, Phys. Lett. B **383**, 345 (1996). E. Malkawi, T. Tait, and C.P. Yuan, Phys. Lett. B **385**, 304 (1996).
12. A. Donini *et al.*, Nucl. Phys. B **507**, 51 (1997).
13. R.S. Chivukula *et al.*, Phys. Rev. D **74**, 075011 (2006). H.J. He, T. Tait, and C.P. Yuan, Phys. Rev. D **62**, 011702 (2000).
14. F. Pisano and V. Pleitez, Phys. Rev. D **46**, 410 (1992); **51**, 3865 (1995).
15. Z. Sullivan, Phys. Rev. D **66**, 075011 (2002).
16. G. Aad *et al.* [ATLAS Collab.], JHEP **1409**, 037 (2014).
17. T.G. Rizzo, JHEP **0705**, 037 (2007); E. Boos *et al.*, Phys. Lett. B **655**, 245 (2007).
18. V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. D **91**, 092005 (2015).
19. CMS Collab., note PAS-EXO-15-006, Dec. 2015.
20. ATLAS Collab., note CONF-2015-063, Dec. 2015.
21. F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **74**, 2900 (1995); S. Abachi *et al.* [D0 Collab.], Phys. Lett. B **358**, 405 (1995).
22. F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **67**, 2609 (1991).
23. V. Khachatryan *et al.* [CMS Collab.], Phys. Lett. B **755**, 196 (2016).
24. G. Aad *et al.* [ATLAS Collab.], JHEP **1507**, 162 (2015).
25. V. Khachatryan *et al.* [CMS Collab.], Eur. Phys. J. C **74**, 3149 (2014).
26. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. B **743**, 235 (2015); Eur. Phys. J. C **75**, 165 (2015).
27. V. Khachatryan *et al.* [CMS Collab.], JHEP **1602**, 122 (2016); JHEP **1405**, 108 (2014).
28. CMS Collab., note PAS-B2G-15-004, Dec. 2015.
29. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. Lett. **115**, 061801 (2015).
30. T.M.P. Tait, C.-P. Yuan, Phys. Rev. D **63**, 014018 (2000).
31. J. Alitti *et al.* [UA2 Collab.], Nucl. Phys. B **400**, 3 (1993).
32. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. D **79**, 112002 (2009).
33. CMS Collab., note PAS-EXO-14-005, Oct. 2015.
34. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. B **754**, 302 (2016); Phys. Rev. D **91**, 052007 (2015); V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. Lett. **116**, 071801 (2016) Phys. Rev. D **91**, 052009 (2015).
35. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. B **737**, 223 (2014); Eur. Phys. J. C **75**, 69 (2015); Eur. Phys. J. C **75**, 209 (2015) [Eur. Phys. J. C **75**, 370 (2015)]; JHEP **1512**, 055 (2015).
36. ATLAS Collab., notes CONF-2015-068; CONF-2015-071; CONF-2015-073; CONF-2015-075, Dec. 2015.
37. V. Khachatryan *et al.* [CMS Collab.], JHEP **1408**, 173 (2014), Phys. Lett. B **740**, 83 (2015).

38. CMS Collab., note PAS-EXO-15-002, Dec. 2015.
39. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. B **755**, 285 (2016).
40. H.J. He *et al.*, Phys. Rev. D **78**, 031701 (2008).
41. G. Aad *et al.* [ATLAS Collab.], Eur. Phys. J. C **75**, 263 (2015).
42. ATLAS Collab., note CONF-2015-074, Dec. 2015.
43. V. Khachatryan *et al.* [CMS Collab.], JHEP **1602**, 145 (2016).
44. D. Pappadopulo *et al.*, JHEP **1409**, 060 (2014).
45. See the particle listings for  $W'$  in this *Review*.
46. Y. Zhang *et al.*, Phys. Rev. D **76**, 091301 (2007); S. Bertolini, A. Maiezza, and F. Nesti, Phys. Rev. D **89**, 095028 (2014).
47. P. Langacker and S.U. Sankar, Phys. Rev. D **40**, 1569 (1989).
48. J. F. Bueno *et al.* [TWIST Collab.], Phys. Rev. D **84**, 032005 (2011).
49. See Fig. 5 of G. Prezeau, M. Ramsey-Musolf, and P. Vogel, Phys. Rev. D **68**, 034016 (2003).

### MASS LIMITS for $W'$ (Heavy Charged Vector Boson Other Than $W$ ) in Hadron Collider Experiments

Couplings of  $W'$  to quarks and leptons are taken to be identical with those of  $W$ . The following limits are obtained from  $p\bar{p}$  or  $pp \rightarrow W'X$  with  $W'$  decaying to the mode indicated in the comments. New decay channels (*e.g.*,  $W' \rightarrow WZ$ ) are assumed to be suppressed. The most recent preliminary results can be found in the “ $W'$ -boson searches” review above.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
none 400–1590	95	1 AAD	15AU ATLS	$W' \rightarrow WZ$
none 1500–1760	95	2 AAD	15AV ATLS	$W' \rightarrow tb$
none 300–1490	95	3 AAD	15AZ ATLS	$W' \rightarrow WZ$
none 1300–1500	95	4 AAD	15CP ATLS	$W' \rightarrow WZ$
none 500–1920	95	5 AAD	15R ATLS	$W' \rightarrow tb$
none 800–2450	95	6 AAD	15V ATLS	$W' \rightarrow q\bar{q}$
>1470	95	7 KHACHATRY..15C	CMS	$W' \rightarrow WZ$
>3710	95	8 KHACHATRY..15T	CMS	$W' \rightarrow \nu\nu, \mu\nu$
none 1200–1900 and 2000–2200	95	9 KHACHATRY..15V	CMS	$W' \rightarrow q\bar{q}$
>3240	95	AAD	14AI ATLS	$W' \rightarrow \nu\nu, \mu\nu$
none 200–1520	95	10 AAD	14S ATLS	$W' \rightarrow WZ$
none 1000–1700	95	11 KHACHATRY..14A	CMS	$W' \rightarrow WZ$
none 1000–3010	95	12 KHACHATRY..14O	CMS	$W' \rightarrow N\ell \rightarrow \ell\ell jj$
none 800–1510	95	13 CHATRCHYAN13E	CMS	$W' \rightarrow tb$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 300–880	95	14 AAD	15BB ATLS	$W' \rightarrow Wh$
		15 AALTONEN	15C CDF	$W' \rightarrow tb$
		16 AAD	14AT ATLS	$W' \rightarrow W\gamma$
		17 KHACHATRY..14A	CMS	$W' \rightarrow WZ$
none 500–950	95	18 AAD	13AO ATLS	$W' \rightarrow WZ$
none 1100–1680	95	AAD	13D ATLS	$W' \rightarrow q\bar{q}$
none 1000–1920	95	CHATRCHYAN13A	CMS	$W' \rightarrow q\bar{q}$
		19 CHATRCHYAN13AJ	CMS	$W' \rightarrow WZ$
>2900	95	20 CHATRCHYAN13AQ	CMS	$W' \rightarrow \nu\nu, \mu\nu$
none 700–940	95	21 CHATRCHYAN13U	CMS	$W' \rightarrow WZ$
none 700–1130	95	22 AAD	12AV ATLS	$W' \rightarrow tb$
none 200–760	95	23 AAD	12BB ATLS	$W' \rightarrow WZ$
		24 AAD	12CK ATLS	$W' \rightarrow \bar{t}q$
>2550	95	25 AAD	12CR ATLS	$W' \rightarrow \nu\nu, \mu\nu$
		26 AAD	12M ATLS	$W' \rightarrow N\ell \rightarrow \ell\ell jj$
		27 AALTONEN	12N CDF	$W' \rightarrow \bar{t}q$
none 200–1143	95	23 CHATRCHYAN12AF	CMS	$W' \rightarrow WZ$
		28 CHATRCHYAN12AR	CMS	$W' \rightarrow \bar{t}q$
		29 CHATRCHYAN12BG	CMS	$W' \rightarrow N\ell \rightarrow \ell\ell jj$
>1120	95	AALTONEN	11C CDF	$W' \rightarrow \nu\nu$
none 180–690	95	30 ABAZOV	11H D0	$W' \rightarrow WZ$
none 600–863	95	31 ABAZOV	11L D0	$W' \rightarrow tb$
none 285–516	95	32 AALTONEN	10N CDF	$W' \rightarrow WZ$
none 280–840	95	33 AALTONEN	09AC CDF	$W' \rightarrow q\bar{q}$
>1000	95	ABAZOV	08C D0	$W' \rightarrow \nu\nu$
none 300–800	95	ABAZOV	04C D0	$W' \rightarrow q\bar{q}$
none 225–536	95	34 ACOSTA	03B CDF	$W' \rightarrow tb$
none 200–480	95	35 AFFOLDER	02C CDF	$W' \rightarrow WZ$
> 786	95	36 AFFOLDER	01I CDF	$W' \rightarrow \nu\nu, \mu\nu$
none 300–420	95	37 ABE	97G CDF	$W' \rightarrow q\bar{q}$
> 720	95	38 ABACHI	96C D0	$W' \rightarrow \nu\nu$
> 610	95	39 ABACHI	95E D0	$W' \rightarrow \nu\nu, \tau\nu$
none 260–600	95	40 RIZZO	93 RVUE	$W' \rightarrow q\bar{q}$

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## New Heavy Bosons

- <sup>1</sup> AAD 15Au search for  $W'$  decaying into the  $WZ$  final state with  $W \rightarrow q\bar{q}'$ ,  $Z \rightarrow \ell^+\ell^-$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>2</sup> AAD 15Av limit is for a SM like right-handed  $W'$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV.  $W' \rightarrow \ell\nu$  decay is assumed to be forbidden.
- <sup>3</sup> AAD 15Az search for  $W'$  decaying into the  $WZ$  final state with  $W \rightarrow \ell\nu$ ,  $Z \rightarrow q\bar{q}$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>4</sup> AAD 15CP search for  $W'$  decaying into the  $WZ$  final state with  $W \rightarrow q\bar{q}$ ,  $Z \rightarrow q\bar{q}$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>5</sup> AAD 15R limit is for a SM like right-handed  $W'$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV.  $W' \rightarrow \ell\nu$  decay is assumed to be forbidden.
- <sup>6</sup> AAD 15v search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>7</sup> KHACHATRYAN 15c search for  $W'$  decaying via  $WZ$  to fully leptonic final states using  $pp$  collisions at  $\sqrt{s}=8$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = M_W/M_{W'}$ .
- <sup>8</sup> KHACHATRYAN 15T limit is for  $W'$  with SM-like coupling which interferes the SM  $W$  boson constructively using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. For  $W'$  without interference, the limit becomes  $> 3280$  GeV.
- <sup>9</sup> KHACHATRYAN 15v search new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>10</sup> AAD 14s search for  $W'$  decaying into the  $WZ$  final state with  $W \rightarrow \ell\nu$ ,  $Z \rightarrow \ell\ell$  using  $pp$  collisions at  $\sqrt{s}=8$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>11</sup> KHACHATRYAN 14 search for  $W'$  decaying into  $WZ$  final state with  $W \rightarrow q\bar{q}$ ,  $Z \rightarrow q\bar{q}$  using  $pp$  collisions at  $\sqrt{s}=8$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>12</sup> KHACHATRYAN 14o search for right-handed  $W_R$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.  $W_R$  is assumed to decay into  $\ell$  and hypothetical heavy neutrino  $N$ , with  $N$  decaying into  $\ell jj$ . The quoted limit is for  $M_{\nu_{eR}} = M_{\nu_{\mu R}} = M_{W_R}/2$ . See their Fig. 3 and Fig. 5 for excluded regions in the  $M_{W_R} - M_{\nu}$  plane.
- <sup>13</sup> CHATRCHYAN 13E limit is for  $W'$  with SM-like coupling which interferes with the SM  $W$  boson using  $pp$  collisions at  $\sqrt{s}=7$  TeV. For  $W'$  with right-handed coupling, the bound becomes  $>1850$  GeV ( $>1910$  GeV) if  $W'$  decays to both leptons and quarks (only to quarks). If both left- and right-handed couplings are present, the limit becomes  $>1640$  GeV.
- <sup>14</sup> AAD 15BB search for  $W'$  decaying into  $Wh$  with  $W \rightarrow \ell\nu$ ,  $h \rightarrow b\bar{b}$ . See their Fig. 4 for the exclusion limits in the heavy vector triplet benchmark model parameter space.
- <sup>15</sup> AALTONEN 15C limit is for a SM-like right-handed  $W'$  assuming  $W' \rightarrow \ell\nu$  decays are forbidden, using  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV. See their Fig. 3 for limit on  $g_{W'WZ}$ .
- <sup>16</sup> AAD 14AT search for a narrow charged vector boson decaying to  $W\gamma$ . See their Fig. 3a for the exclusion limit in  $M_{W'} - \sigma_B$  plane.
- <sup>17</sup> KHACHATRYAN 14A search for  $W'$  decaying into the  $WZ$  final state with  $W \rightarrow \ell\nu$ ,  $Z \rightarrow q\bar{q}$ , or  $W \rightarrow q\bar{q}$ ,  $Z \rightarrow \ell\ell$ .  $pp$  collisions data at  $\sqrt{s}=8$  TeV are used for the search. See their Fig. 13 for the exclusion limit on the number of events in the mass-width plane.
- <sup>18</sup> AAD 13Ao search for  $W'$  decaying into the  $WZ$  final state with  $W \rightarrow \ell\nu$ ,  $Z \rightarrow 2j$  using  $pp$  collisions at  $\sqrt{s}=7$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>19</sup> CHATRCHYAN 13AJ search for resonances decaying to  $WZ$  pair, using the hadronic decay modes of  $W$  and  $Z$ , in  $pp$  collisions at  $\sqrt{s}=7$  TeV. See their Fig. 7 for the limit on the cross section.
- <sup>20</sup> CHATRCHYAN 13AQ limit is for  $W'$  with SM-like coupling which interferes with the SM  $W$  boson using  $pp$  collisions at  $\sqrt{s}=7$  TeV.
- <sup>21</sup> CHATRCHYAN 13U search for  $W'$  decaying to the  $WZ$  final state, with  $W$  decaying into jets, in  $pp$  collisions at  $\sqrt{s}=7$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>22</sup> The AAD 12Av quoted limit is for a SM-like right-handed  $W'$  using  $pp$  collisions at  $\sqrt{s}=7$  TeV.  $W' \rightarrow \ell\nu$  decay is assumed to be forbidden.
- <sup>23</sup> AAD 12BB use  $pp$  collisions data at  $\sqrt{s}=7$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ .
- <sup>24</sup> AAD 12CK search for  $pp \rightarrow tW'$ ,  $W' \rightarrow \tau q$  events in  $pp$  collisions. See their Fig. 5 for the limit on  $\sigma \cdot B$ .
- <sup>25</sup> AAD 12CR use  $pp$  collisions at  $\sqrt{s}=7$  TeV.
- <sup>26</sup> AAD 12M search for right-handed  $W_R$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.  $W_R$  is assumed to decay into  $\ell$  and hypothetical heavy neutrino  $N$ , with  $N$  decaying into  $\ell jj$ . See their Fig. 4 for the limit in the  $m_N - m_{W'}$  plane.
- <sup>27</sup> AALTONEN 12N search for  $p\bar{p} \rightarrow tW'$ ,  $W' \rightarrow \tau d$  events in  $p\bar{p}$  collisions. See their Fig. 3 for the limit on  $\sigma \cdot B$ .
- <sup>28</sup> CHATRCHYAN 12AR search for  $pp \rightarrow tW'$ ,  $W' \rightarrow \tau d$  events in  $pp$  collisions. See their Fig. 2 for the limit on  $\sigma \cdot B$ .
- <sup>29</sup> CHATRCHYAN 12BG search for right-handed  $W_R$  in  $pp$  collisions  $\sqrt{s} = 7$  TeV.  $W_R$  is assumed to decay into  $\ell$  and hypothetical heavy neutrino  $N$ , with  $N$  decaying into  $\ell jj$ . See their Fig. 3 for the limit in the  $m_N - m_{W'}$  plane.
- <sup>30</sup> ABZOV 11H use data from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV. The quoted limit is obtained assuming  $W'WZ$  coupling strength is the same as the ordinary  $WWZ$  coupling strength in the Standard Model.
- <sup>31</sup> ABZOV 11L limit is for  $W'$  with SM-like coupling which interferes with the SM  $W$  boson, using  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV. For  $W'$  with right-handed coupling, the bound becomes  $>885$  GeV ( $>890$  GeV) if  $W'$  decays to both leptons and quarks (only to quarks). If both left- and right-handed couplings present, the limit becomes  $>916$  GeV.
- <sup>32</sup> AALTONEN 10N use  $p\bar{p}$  collision data at  $\sqrt{s}=1.96$  TeV. The quoted limit assumes  $g_{W'WZ}/g_{WWZ} = (M_W/M_{W'})^2$ . See their Fig. 4 for limits in mass-coupling plane.

- <sup>33</sup> AALTONEN 09AC search for new particle decaying to dijets using  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV.
- <sup>34</sup> The ACOSTA 03B quoted limit is for  $M_{W'} \gg M_{\nu_R}$ , using  $p\bar{p}$  collisions at  $\sqrt{s}=1.8$  TeV. For  $M_{W'} < M_{\nu_R}$ ,  $M_{W'}$  between 225 and 566 GeV is excluded.
- <sup>35</sup> The quoted limit is obtained assuming  $W'WZ$  coupling strength is the same as the ordinary  $WWZ$  coupling strength in the Standard Model, using  $p\bar{p}$  collisions at  $\sqrt{s}=1.8$  TeV. See their Fig. 2 for the limits on the production cross sections as a function of the  $W'$  width.
- <sup>36</sup> AFFOLDER 01i combine a new bound on  $W' \rightarrow e\nu$  of 754 GeV, using  $p\bar{p}$  collisions at  $\sqrt{s}=1.8$  TeV, with the bound of ABE 00 on  $W' \rightarrow \mu\nu$  to obtain quoted bound.
- <sup>37</sup> ABE 97c search for new particle decaying to dijets using  $p\bar{p}$  collisions at  $\sqrt{s}=1.8$  TeV.
- <sup>38</sup> For bounds on  $W_R$  with nonzero right-handed mass, see Fig. 5 from ABACHI 96c.
- <sup>39</sup> ABACHI 95E assume that the decay  $W' \rightarrow WZ$  is suppressed and that the neutrino from  $W'$  decay is stable and has a mass significantly less  $m_{W'}$ .
- <sup>40</sup> RIZZO 93 analyses CDF limit on possible two-jet resonances. The limit is sensitive to the inclusion of the assumed  $K$  factor.

### $W_R$ (Right-Handed $W$ Boson) MASS LIMITS

Assuming a light right-handed neutrino, except for BEALL 82, LANGACKER 89b, and COLANGELO 91.  $g_R = g_L$  assumed. [Limits in the section MASS LIMITS for  $W'$  below are also valid for  $W_R$  if  $m_{\nu_R} \ll m_{W_R}$ .] Some limits assume manifest left-right symmetry, i.e., the equality of left- and right Cabibbo-Kobayashi-Maskawa matrices. For a comprehensive review, see LANGACKER 89b. Limits on the  $W_L - W_R$  mixing angle  $\zeta$  are found in the next section. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 592	90	1 BUENO	11 TWST	$\mu$ decay
> 715	90	2 CZAKON	99 RVUE	Electroweak
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 235	90	3 PRIEELS	14 PIE3	$\mu$ decay
> 245	90	4 WAUTERS	10 CNTR	$^{60}\text{Co}$ $\beta$ decay
>2500	90	5 ZHANG	08 THEO	$m_{K_0^0} - m_{K_S^0}$
> 180	90	6 MELCONIAN	07 CNTR	$^{37}\text{K}$ $\beta^+$ decay
> 290.7	90	7 SCHUMANN	07 CNTR	Polarized neutron decay
[> 3300]	95	8 CYBURT	05 COSM	Nucleosynthesis; light $\nu_R$
> 310	90	9 THOMAS	01 CNTR	$\beta^+$ decay
> 137	95	10 ACKERSTAFF	99D OPAL	$\tau$ decay
>1400	68	11 BARENBOIM	98 RVUE	Electroweak, $Z-Z'$ mixing
> 549	68	12 BARENBOIM	97 RVUE	$\mu$ decay
> 220	95	13 STAHL	97 RVUE	$\tau$ decay
> 220	90	14 ALLET	96 CNTR	$\beta^+$ decay
> 281	90	15 KUZNETSOV	95 CNTR	Polarized neutron decay
> 282	90	16 KUZNETSOV	94b CNTR	Polarized neutron decay
> 439	90	17 BHATTACH...	93 RVUE	$Z-Z'$ mixing
> 250	90	18 SEVERIJNS	93 CNTR	$\beta^+$ decay
> 475	90	19 IMAZATO	92 CNTR	$K^+$ decay
> 240	90	20 POLAK	92b RVUE	$\mu$ decay
> 496	90	21 AQUINO	91 RVUE	Neutron decay
> 700	90	22 COLANGELO	91 THEO	Neutron and muon decay
> 477	90	23 POLAK	91 RVUE	$m_{K_0^0} - m_{K_S^0}$
[none 540-23000]	90	24 BARBIERI	89b ASTR	$\mu$ decay
> 300	90	25 LANGACKER	89b RVUE	SN 1987A; light $\nu_R$
> 160	90	26 BALKE	88 CNTR	General
> 406	90	27 JODIDIO	86 ELEC	$\mu \rightarrow e\nu\bar{\nu}$
> 482	90	28 JODIDIO	86 ELEC	Any $\zeta$
> 800	95	29 MOHAPATRA	86 RVUE	$\zeta = 0$
> 400	95	28 STOKER	85 ELEC	$\text{SU}(2)_L \times \text{SU}(2)_R \times \text{U}(1)$
> 475	95	28 STOKER	85 ELEC	Any $\zeta$
> 380	90	29 BERGSMAN	83 CHRM	$\zeta < 0.041$
> 1600	90	30 CARR	83 ELEC	$\nu_\mu e \rightarrow \mu\nu_e$
		31 BEALL	82 THEO	$\mu^+$ decay
				$m_{K_L^0} - m_{K_S^0}$

- <sup>1</sup> The quoted limit is for manifest left-right symmetric model.
- <sup>2</sup> CZAKON 99 perform a simultaneous fit to charged and neutral sectors.
- <sup>3</sup> PRIEELS 14 limit is from  $\mu^+ \rightarrow e^+ \nu \bar{\nu}$  decay parameter  $\xi''$ , which is determined by the positron polarization measurement.
- <sup>4</sup> WAUTERS 10 limit is from a measurement of the asymmetry parameter of polarized  $^{60}\text{Co}$   $\beta$  decays. The listed limit assumes no mixing.
- <sup>5</sup> ZHANG 08 limit uses a lattice QCD calculation of the relevant hadronic matrix elements, while BEALL 82 limit used the vacuum saturation approximation.
- <sup>6</sup> MELCONIAN 07 measure the neutrino angular asymmetry in  $\beta^+$ -decays of polarized  $^{37}\text{K}$ , stored in a magneto-optical trap. Result is consistent with SM prediction and does not constrain the  $W_L - W_R$  mixing angle appreciably.
- <sup>7</sup> SCHUMANN 07 limit is from measurements of the asymmetry  $\langle \vec{p}_\nu \cdot \vec{\sigma}_N \rangle$  in the  $\beta$  decay of polarized neutrons. Zero mixing is assumed.
- <sup>8</sup> CYBURT 05 limit follows by requiring that three light  $\nu_R$ 's decouple when  $T_{dec} > 140$  MeV. For different  $T_{dec}$ , the bound becomes  $M_{W_R} > 3.3 \text{ TeV } (T_{dec} / 140 \text{ MeV})^{3/4}$ .
- <sup>9</sup> THOMAS 01 limit is from measurement of  $\beta^+$  polarization in decay of polarized  $^{12}\text{N}$ . The listed limit assumes no mixing.
- <sup>10</sup> ACKERSTAFF 99D limit is from  $\tau$  decay parameters. Limit increase to 145 GeV for zero mixing.
- <sup>11</sup> BARENBOIM 98 assumes minimal left-right model with Higgs of  $\text{SU}(2)_R$  in  $\text{SU}(2)_L$  doublet. For Higgs in  $\text{SU}(2)_L$  triplet,  $m_{W_R} > 1100$  GeV. Bound calculated from effect of corresponding  $Z_{LR}$  on electroweak data through  $Z-Z_{LR}$  mixing.

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

- <sup>12</sup> The quoted limit is from  $\mu$  decay parameters. BARENBOIM 97 also evaluate limit from  $K_L - K_S$  mass difference.
- <sup>13</sup> STAHL 97 limit is from fit to  $\tau$ -decay parameters.
- <sup>14</sup> ALLET 96 measured polarization-asymmetry correlation in  $^{12}\text{N} \beta^+$  decay. The listed limit assumes zero  $L$ - $R$  mixing.
- <sup>15</sup> KUZNETSOV 95 limit is from measurements of the asymmetry  $\langle \bar{p}_\nu \cdot \sigma_n \rangle$  in the  $\beta$  decay of polarized neutrons. Zero mixing assumed. See also KUZNETSOV 94b.
- <sup>16</sup> KUZNETSOV 94b limit is from measurements of the asymmetry  $\langle \bar{p}_\nu \cdot \sigma_n \rangle$  in the  $\beta$  decay of polarized neutrons. Zero mixing assumed.
- <sup>17</sup> BHATTACHARYYA 93 uses  $Z$ - $Z'$  mixing limit from LEP '90 data, assuming a specific Higgs sector of  $\text{SU}(2)_L \times \text{SU}(2)_R \times \text{U}(1)$  gauge model. The limit is for  $m_t = 200$  GeV and slightly improves for smaller  $m_t$ .
- <sup>18</sup> SEVERIJNS 93 measured polarization-asymmetry correlation in  $^{107}\text{In} \beta^+$  decay. The listed limit assumes zero  $L$ - $R$  mixing. Value quoted here is from SEVERIJNS 94 erratum.
- <sup>19</sup> IMAZATO 92 measure positron asymmetry in  $K^+ \rightarrow \mu^+ \nu_\mu$  decay and obtain  $\xi P_\mu > 0.990$  (90% CL). If  $W_R$  couples to  $u\bar{s}$  with full weak strength ( $V_{us}^R = 1$ ), the result corresponds to  $m_{W_R} > 653$  GeV. See their Fig. 4 for  $m_{W_R}$  limits for general  $|V_{us}^R|^2 = 1 - |V_{ud}^R|^2$ .
- <sup>20</sup> POLAK 92b limit is from fit to muon decay parameters and is essentially determined by JODIDIO 86 data assuming  $\zeta = 0$ . Supersedes POLAK 91.
- <sup>21</sup> AQUINO 91 limits obtained from neutron lifetime and asymmetries together with unitarity of the CKM matrix. Manifest left-right symmetry assumed. Stronger of the two limits also includes muon decay results.
- <sup>22</sup> COLANGELO 91 limit uses hadronic matrix elements evaluated by QCD sum rule and is less restrictive than BEALL 82 limit which uses vacuum saturation approximation. Manifest left-right symmetry assumed.
- <sup>23</sup> POLAK 91 limit is from fit to muon decay parameters and is essentially determined by JODIDIO 86 data assuming  $\zeta = 0$ . Superseded by POLAK 92b.
- <sup>24</sup> BARBIERI 89b limit holds for  $m_{\nu_R} \leq 10$  MeV.
- <sup>25</sup> LANGACKER 89b limit is for any  $\nu_R$  mass (either Dirac or Majorana) and for a general class of right-handed quark mixing matrices.
- <sup>26</sup> BALKE 88 limit is for  $m_{\nu_R} = 0$  and  $m_{\nu_{\mu R}} \leq 50$  MeV. Limits come from precise measurements of the muon decay asymmetry as a function of the positron energy.
- <sup>27</sup> JODIDIO 86 is the same TRIUMF experiment as STOKER 85 (and CARR 83); however, it uses a different technique. The results given here are combined results of the two techniques. The technique here involves precise measurement of the end-point  $e^+$  spectrum in the decay of the highly polarized  $\mu^+$ .
- <sup>28</sup> STOKER 85 is same TRIUMF experiment as CARR 83. Here they measure the decay  $e^+$  spectrum asymmetry above 46 MeV/c using a muon-spin-rotation technique. Assumed a light right-handed neutrino. Quoted limits are from combining with CARR 83.
- <sup>29</sup> BERGSMAN 83 set limit  $m_{W_2}/m_{W_1} > 1.9$  at CL = 90%.
- <sup>30</sup> CARR 83 is TRIUMF experiment with a highly polarized  $\mu^+$  beam. Looked for deviation from  $V-A$  at the high momentum end of the decay  $e^+$  energy spectrum. Limit from previous world-average muon polarization parameter is  $m_{W_R} > 240$  GeV. Assumes a light right-handed neutrino.
- <sup>31</sup> BEALL 82 limit is obtained assuming that  $W_R$  contribution to  $K_L^0 - K_S^0$  mass difference is smaller than the standard one, neglecting the top quark contributions. Manifest left-right symmetry assumed.

### Limit on $W_L$ - $W_R$ Mixing Angle $\zeta$

Lighter mass eigenstate  $W_1 = W_L \cos \zeta - W_R \sin \zeta$ . Light  $\nu_R$  assumed unless noted. Values in brackets are from cosmological and astrophysical considerations.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.020 to 0.017	90	BUENO 11	TWST	$\mu \rightarrow e \nu \bar{\nu}$
< 0.022	90	MACDONALD 08	TWST	$\mu \rightarrow e \nu \bar{\nu}$
< 0.12	95	1 ACKERSTAFF 99b	OPAL	$\tau$ decay
< 0.013	90	2 CZAKON 99	RVUE	Electroweak
< 0.0333		3 BARENBOIM 97	RVUE	$\mu$ decay
< 0.04	90	4 MISHRA 92	CCFR	$\nu N$ scattering
-0.0006 to 0.0028	90	5 AQUINO 91	RVUE	
[none 0.00001-0.02]		6 BARBIERI 89b	ASTR	SN 1987A
< 0.040	90	7 JODIDIO 86	ELEC	$\mu$ decay
-0.056 to 0.040	90	7 JODIDIO 86	ELEC	$\mu$ decay

- <sup>1</sup> ACKERSTAFF 99b limit is from  $\tau$  decay parameters.
- <sup>2</sup> CZAKON 99 perform a simultaneous fit to charged and neutral sectors.
- <sup>3</sup> The quoted limit is from  $\mu$  decay parameters. BARENBOIM 97 also evaluate limit from  $K_L - K_S$  mass difference.
- <sup>4</sup> MISHRA 92 limit is from the absence of extra large- $x$ , large- $y$   $\bar{\nu}_\mu N \rightarrow \bar{\nu}_\mu X$  events at Tevatron, assuming left-handed  $\nu$  and right-handed  $\bar{\nu}$  in the neutrino mass. The result gives  $\zeta^2(1 - 2m_{W_1}^2/m_{W_2}^2) < 0.0015$ . The limit is independent of  $\nu_R$  mass.
- <sup>5</sup> AQUINO 91 limits obtained from neutron lifetime and asymmetries together with unitarity of the CKM matrix. Manifest left-right asymmetry is assumed.
- <sup>6</sup> BARBIERI 89b limit holds for  $m_{\nu_R} \leq 10$  MeV.
- <sup>7</sup> First JODIDIO 86 result assumes  $m_{W_R} = \infty$ , second is for unconstrained  $m_{W_R}$ .

### Z'-BOSON SEARCHES

Revised Jan. 2016 by M.-C. Chen (UC Irvine), B.A. Dobrescu (Fermilab) and S. Willocq (Univ. of Massachusetts).

The  $Z'$  boson is a massive, electrically-neutral and color-singlet hypothetical particle of spin 1. This particle is predicted

in many extensions of the Standard Model (SM) and has been the object of extensive phenomenological studies [1].

**$Z'$  boson couplings to quarks and leptons.** The couplings of a  $Z'$  boson to the first-generation fermions are given by

$$Z'_\mu (g_u^L \bar{u}_L \gamma^\mu u_L + g_d^L \bar{d}_L \gamma^\mu d_L + g_u^R \bar{u}_R \gamma^\mu u_R + g_d^R \bar{d}_R \gamma^\mu d_R + g_\nu^L \bar{\nu}_L \gamma^\mu \nu_L + g_e^L \bar{e}_L \gamma^\mu e_L + g_e^R \bar{e}_R \gamma^\mu e_R), \quad (1)$$

where  $u, d, \nu$  and  $e$  are the quark and lepton fields in the mass eigenstate basis, and the coefficients  $g_u^L, g_d^L, g_u^R, g_d^R, g_\nu^L, g_e^L, g_e^R$  are real dimensionless parameters. If the  $Z'$  couplings to quarks and leptons are generation-independent, then these seven parameters describe the couplings of the  $Z'$  boson to all SM fermions. More generally, however, the  $Z'$  couplings to fermions are generation-dependent, in which case Eq. (1) may be written with generation indices  $i, j = 1, 2, 3$  labeling the quark and lepton fields, and with the seven coefficients promoted to  $3 \times 3$  Hermitian matrices (e.g.,  $g_{eij}^L \bar{e}_L^i \gamma^\mu e_L^j$ , where  $e_L^j$  is the left-handed muon, etc.).

These parameters describing the  $Z'$  boson interactions with quarks and leptons are subject to some theoretical constraints. Quantum field theories that include a heavy spin-1 particle are well behaved at high energies only if that particle is a gauge boson associated with a spontaneously broken gauge symmetry. Quantum effects preserve the gauge symmetry only if the couplings of the gauge boson to fermions satisfy anomaly cancellation conditions. Furthermore, the fermion charges under the new gauge symmetry are constrained by the requirement that the quarks and leptons get masses from gauge-invariant interactions with Higgs fields.

The relation between the couplings displayed in Eq. (1) and the gauge charges  $z_{fi}^L$  and  $z_{fi}^R$  of the fermions  $f = u, d, \nu, e$  involves the unitary  $3 \times 3$  matrices  $V_f^L$  and  $V_f^R$  that transform the gauge eigenstate fermions  $f_L^i$  and  $f_R^i$ , respectively, into the mass eigenstates. The  $Z'$  couplings are also modified if the new gauge boson in the gauge eigenstate basis ( $\tilde{Z}'_\mu$ ) has a kinetic mixing  $(-\chi/2)B^{\mu\nu}\tilde{Z}'_{\mu\nu}$  with the hypercharge gauge boson  $B^\mu$  ( $\chi$  is a dimensionless parameter), or a mass mixing  $\delta M^2 \tilde{Z}^\mu \tilde{Z}'_\mu$  with the linear combination ( $\tilde{Z}_\mu$ ) of neutral bosons that couples as the SM  $Z$  boson [2]. Since both the kinetic and mass mixings shift the mass and couplings of the  $Z$  boson, electroweak measurements impose upper limits on  $\chi$  and  $\delta M^2/(M_{Z'}^2 - M_Z^2)$  of the order of  $10^{-3}$  [3]. Keeping only linear terms in these two small quantities, the couplings of the mass-eigenstate  $Z'$  boson are given by

$$g_{fij}^L = g_z V_{fii'}^L z_{f'i'}^L (V_f^L)_{ij}^\dagger + \frac{e}{c_W} \left( \frac{s_W \chi M_{Z'}^2 + \delta M^2}{2s_W (M_{Z'}^2 - M_Z^2)} \sigma_f^3 - \epsilon Q_f \right),$$

$$g_{fij}^R = g_z V_{fii'}^R z_{f'i'}^R (V_f^R)_{ij}^\dagger - \frac{e}{c_W} \epsilon Q_f, \quad (2)$$

where  $g_z$  is the new gauge coupling,  $Q_f$  is the electric charge of  $f$ ,  $e$  is the electromagnetic gauge coupling,  $s_W$  and  $c_W$  are the

**Table 1:** Examples of generation-independent  $U(1)'$  charges for quarks and leptons. The parameter  $x$  is an arbitrary rational number. Anomaly cancellation requires certain new fermions [5].

fermion	$U(1)_{B-xL}$	$U(1)_{10+5}$	$U(1)_{d-xu}$	$U(1)_{q+xu}$
$(u_L, d_L)$	1/3	1/3	0	1/3
$u_R$	1/3	-1/3	$-x/3$	$x/3$
$d_R$	1/3	$-x/3$	1/3	$(2-x)/3$
$(\nu_L, e_L)$	$-x$	$x/3$	$(-1+x)/3$	-1
$e_R$	$-x$	-1/3	$x/3$	$-(2+x)/3$

sine and cosine of the weak mixing angle,  $\sigma_f^3 = +1$  for  $f = u, \nu$  and  $\sigma_f^3 = -1$  for  $f = d, e$ , and

$$\epsilon = \frac{\chi (M_{Z'}^2 - c_W^2 M_Z^2) + s_W \delta M^2}{M_{Z'}^2 - M_Z^2}. \quad (3)$$

The interaction of the  $Z'$  boson with a pair of  $W$  bosons has the form

$$\left( i(W_\mu^- Z'_\nu - W_\nu^- Z'_\mu) \partial^\mu W^{+\nu} + \text{H.c.} \right) + i(W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) \partial^\mu Z'^{\nu} \quad (4)$$

with a coefficient of order  $M_W^2/M_{Z'}^2$  [4]. The  $Z'$  also couples to one SM Higgs boson and one  $Z$  boson,  $Z'_\mu Z^\mu h^0$ , with a coefficient of order  $M_Z$ .

**$U(1)$  gauge groups.** A simple origin of a  $Z'$  boson is a new  $U(1)'$  gauge symmetry. In that case, the matricial equalities  $z_u^L = z_d^L$  and  $z_\nu^L = z_e^L$  are required by the SM  $SU(2)_W$  gauge symmetry. Given that the  $U(1)'$  interaction is not asymptotically free, the theory may be well-behaved at high energies (*e.g.*, by embedding  $U(1)'$  in a non-Abelian gauge group) only if the charges are commensurate numbers, *i.e.* any ratio of charges is a rational number. Satisfying the anomaly cancellation conditions (which include an equation cubic in charges) with rational numbers is highly nontrivial and in general new fermions charged under  $U(1)'$  are necessary.

Consider first generation-independent couplings (the  $V_f$  matrices then disappear from Eq. (2)) and neglect the  $\tilde{Z} - \tilde{Z}'$  mixing, so that there are five commensurate couplings:  $g_q^L, g_u^R, g_d^R, g_l^L, g_e^R$ . Four sets of charges are displayed in Table 1, each of them spanned by a free parameter  $x$  [5]. The first set, labelled  $B - xL$ , has charges proportional to the baryon number minus  $x$  times the lepton number. These charges allow all SM Yukawa couplings to a Higgs doublet which is neutral under  $U(1)_{B-xL}$ , so that there is no tree-level  $\tilde{Z} - \tilde{Z}'$  mixing. For  $x = 1$  one recovers the  $U(1)_{B-L}$  group, which is non-anomalous in the presence of one “right-handed neutrino” (a chiral fermion that is a singlet under the SM gauge group) per generation. For  $x \neq 1$ , it is necessary to include some fermions that are vector-like (*i.e.* their mass terms are gauge invariant) with respect to the electroweak gauge group and chiral with respect to  $U(1)_{B-xL}$ . In the particular cases  $x = 0$  or  $x \gg 1$ , the  $Z'$  is leptophobic or quark-phobic, respectively.

The second set,  $U(1)_{10+5}$ , has charges that commute with the representations of the  $SU(5)$  grand unified group. Here  $x$  is related to the mixing angle between the two  $U(1)$  bosons encountered in the  $E_6 \rightarrow SU(5) \times U(1) \times U(1)$  symmetry breaking patterns of grand unified theories [1,6]. This set leads to  $\tilde{Z} - \tilde{Z}'$  mass mixing at tree level, such that for a  $Z'$  mass close to the electroweak scale, the measurements at the  $Z$ -pole require some fine tuning between the charges and VEVs of the two Higgs doublets. Vector-like fermions charged under the electroweak gauge group and also carrying color are required (except for  $x = -3$ ) to make this set anomaly free. The particular cases  $x = -3, 1, -1/2$  are usually labelled  $U(1)_\chi$ ,  $U(1)_\psi$ , and  $U(1)_\eta$ , respectively. Under the third set,  $U(1)_{d-xu}$ , the weak-doublet quarks are neutral, and the ratio of  $u_R$  and  $d_R$  charges is  $-x$ . For  $x = 1$  this is the “right-handed” group  $U(1)_R$ . For  $x = 0$ , the charges are those of the  $E_6$ -inspired  $U(1)_I$  group, which requires new quarks and leptons. Other generation-independent sets of  $U(1)'$  charges are given in [7].

In the absence of new fermions charged under the SM group, the most general generation-independent charge assignment is  $U(1)_{q+xu}$ , which is a linear combination of hypercharge and  $B - L$ . Many other anomaly-free solutions exist if generation-dependent charges are allowed. An example is  $B - xL_e - yL_\mu + (y - 3)L_\tau$ , with  $x, y$  free parameters. This allows all fermion masses to be generated by Yukawa couplings to a single Higgs doublet, without inducing tree-level flavor-changing neutral current (FCNC) processes. There are also lepton-flavor dependent charges that allow neutrino masses to arise only from operators of high dimensionality [8].

If the  $SU(2)_W$ -doublet quarks have generation-dependent  $U(1)'$  charges, then the mass eigenstate quarks have flavor off-diagonal couplings to the  $Z'$  boson (see Eq. (1), and note that  $V_u^L (V_d^L)^\dagger$  is the CKM matrix). These are severely constrained by measurements of FCNC processes, which in this case are mediated at tree-level by  $Z'$  boson exchange [9]. The constraints are relaxed if the first and second generation charges are the same, although they are increasingly tightened by the measurements of  $B$  meson properties [10]. If only the  $SU(2)_W$ -singlet quarks have generation-dependent  $U(1)'$  charges, there is more freedom in adjusting the flavor off-diagonal couplings because the  $V_{u,d}^R$  matrices are not observable in the SM.

The anomaly cancellation conditions for  $U(1)'$  could be relaxed only if there is an axion with certain dimension-5 couplings to the gauge bosons. However, such a scenario violates unitarity unless the quantum field theory description breaks down at a scale near  $M_{Z'}$  [11].

**Other models.**  $Z'$  bosons may also arise from larger gauge groups. These may extend the electroweak group, as in  $SU(2) \times SU(2) \times U(1)$ , or may embed the electroweak group, as in  $SU(3)_W \times U(1)$  [12]. If the larger group is spontaneously broken down to  $SU(2)_W \times U(1)_Y \times U(1)'$  at a scale  $v_\star \gg M_{Z'}/g_z$ , then the above discussion applies up to corrections of order  $M_{Z'}^2/(g_z v_\star)^2$ . For  $v_\star \sim M_{Z'}/g_z$ , additional gauge bosons have



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masses comparable to  $M_{Z'}$ , including at least a  $W'$  boson [12]. If the larger gauge group breaks together with the electroweak symmetry directly to the electromagnetic  $U(1)_{\text{em}}$ , then the left-handed fermion charges are no longer correlated ( $z_u^L \neq z_d^L$ ,  $z_\nu^L \neq z_e^L$ ) and a  $Z'W^+W^-$  coupling is induced.

If the electroweak gauge bosons propagate in extra dimensions, then their Kaluza-Klein (KK) excitations include a series of  $Z'$  boson pairs. Each of these pairs can be associated with a different  $SU(2) \times U(1)$  gauge group in four dimensions. The properties of the KK particles depend strongly on the extra-dimensional theory [13]. For example, in universal extra dimensions there is a parity that forces all couplings of Eq. (1) to vanish in the case of the lightest KK bosons, while allowing couplings to pairs of fermions involving a SM and a heavy vector-like fermion. There are also 4-dimensional gauge theories (*e.g.* little Higgs with  $T$  parity) with  $Z'$  bosons exhibiting similar properties. By contrast, in a warped extra dimension, the couplings of Eq. (1) may be sizable even when SM fields propagate along the extra dimension.

$Z'$  bosons may also be composite particles. For example, in confining gauge theories [14], the  $\rho$ -like bound state is a spin-1 boson that may be interpreted as arising from a spontaneously broken gauge symmetry [15].

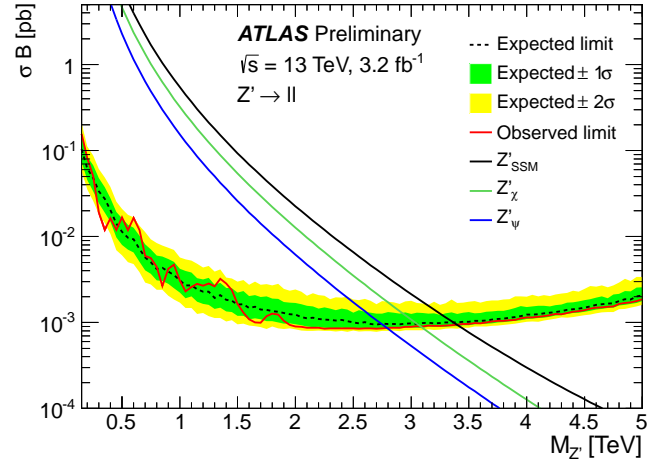
**Resonances versus cascade decays.** In the presence of the couplings shown in Eq. (1), the  $Z'$  boson may be produced in the  $s$ -channel at colliders, and would decay to pairs of fermions. The decay width into a pair of electrons is given by

$$\Gamma(Z' \rightarrow e^+e^-) \simeq \left[ (g_e^L)^2 + (g_e^R)^2 \right] \frac{M_{Z'}}{24\pi}, \quad (5)$$

where small corrections from electroweak loops are not included. The decay width into  $q\bar{q}$  is similar, except for an additional color factor of 3, QCD radiative corrections, and fermion mass corrections. Thus, one may compute the  $Z'$  branching fractions in terms of the couplings of Eq. (1). However, other decay channels, such as  $WW$  or a pair of new particles, could have large widths and need to be added to the total decay width.

As mentioned above, there are theories in which the  $Z'$  couplings are controlled by a discrete symmetry that forbids decays into a pair of SM particles. Typically, such theories involve several new particles, which may be produced only in pairs and undergo cascade decays through  $Z'$  bosons, leading to signals involving some missing (transverse) momentum. Given that the cascade decays depend on the properties of new particles other than the  $Z'$  boson, this case is not discussed further here.

**LEP-II limits.** The  $Z'$  contribution to the cross sections for  $e^+e^- \rightarrow f\bar{f}$  proceeds through an  $s$ -channel  $Z'$  exchange (when  $f = e$ , there are also  $t$ - and  $u$ -channel exchanges). For  $M_{Z'} < \sqrt{s}$ , the  $Z'$  appears as an  $f\bar{f}$  resonance in the radiative return process where photon emission tunes the effective center-of-mass energy to  $M_{Z'}$ . The agreement between the LEP-II measurements and the SM predictions implies that either the  $Z'$  couplings are smaller than or of order  $10^{-2}$ , or else  $M_{Z'}$  is



**Figure 1:** Upper limit on  $\sigma(pp \rightarrow Z'X \rightarrow \ell^+\ell^-X)$  with  $\ell = e$  or  $\mu$  as a function of  $M_{Z'}$  [23], assuming equal couplings for electrons and muons. The lines labelled by  $Z'_\psi$  and  $Z'_\chi$  are theoretical predictions for the  $U(1)_{10+x5}$  models in Table 1 with  $x = -3$  and  $x = +1$ , respectively, for  $g_z$  fixed by an  $E_6$  unification condition. The  $Z'_{\text{SSM}}$  line corresponds to  $Z'$  couplings equal to those of the  $Z$  boson.

above 209 GeV, the maximum energy of LEP-II. In the latter case, the  $Z'$  effects may be approximated up to corrections of order  $s/M_{Z'}^2$  by the contact interactions

$$\frac{g_z^2}{M_{Z'}^2 - s} [\bar{e}\gamma_\mu (z_e^L P_L + z_e^R P_R) e] [\bar{f}\gamma^\mu (z_f^L P_L + z_f^R P_R) f], \quad (6)$$

where  $P_{L,R}$  are chirality projection operators, and the relation between  $Z'$  couplings and charges (see Eq. (2) in the limit where the mass and kinetic mixings are neglected) is used, assuming generation-independent charges. The four LEP collaborations have set limits on the coefficients of such operators for all possible chiral structures and for various combinations of fermions [16]. Thus, one may derive bounds on  $(M_{Z'}/g_z)|z_e^L z_f^L|^{-1/2}$  and the analogous combinations of  $LR$ ,  $RL$  and  $RR$  charges, which are typically on the order of a few TeV. LEP-II limits were derived [5] on the four sets of charges shown in Table 1.

Somewhat stronger bounds can be set on  $M_{Z'}/g_z$  for specific sets of  $Z'$  couplings if the effects of several operators from Eq. (6) are combined. Dedicated analyses by the LEP collaborations have set limits on  $Z'$  bosons for particular values of the gauge coupling (see section 3.5 of [16]).

**Searches at hadron colliders.**  $Z'$  bosons with couplings to quarks (see Eq. (1)) may be produced at hadron colliders in the  $s$ -channel and would show up as resonances in the invariant mass distribution of the decay products. The cross section for



See key on page 601

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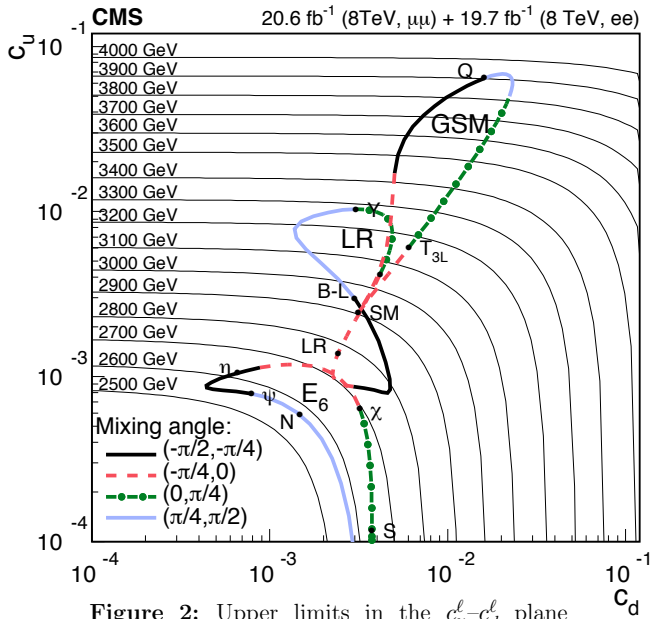
producing a  $Z'$  boson at the LHC, which then decays to some  $f\bar{f}$  final state, takes the form

$$\sigma(pp \rightarrow Z'X \rightarrow f\bar{f}X) \simeq \frac{\pi}{48s} \sum_q c_q^f w_q(s, M_{Z'}^2) \quad (7)$$

for flavor-diagonal couplings to quarks. Here, we have neglected the interference with the SM contribution to  $f\bar{f}$  production, which is a good approximation for a narrow  $Z'$  resonance (deviations from the narrow width approximation are discussed in [17]). The coefficients

$$c_q^f = \left[ (g_q^f)^2 + (g_q^f)^2 \right] B(Z' \rightarrow f\bar{f}) \quad (8)$$

contain all the dependence on the  $Z'$  couplings, while the functions  $w_q$  include all the information about parton distributions and QCD corrections [5,7]. This factorization holds exactly to NLO and the deviations from it induced at NNLO are very small. Note that the  $w_u$  and  $w_d$  functions are substantially larger than the  $w_q$  functions for the other quarks. Eq. (7) also applies to the Tevatron, except for changing the  $pp$  initial state to  $p\bar{p}$ , which implies that the  $w_q(s, M_{Z'}^2)$  functions are replaced by some other functions  $\bar{w}_q((1.96 \text{ TeV})^2, M_{Z'}^2)$ .



**Figure 2:** Upper limits in the  $c_u^\ell - c_d^\ell$  plane ( $\ell = e$  or  $\mu$ ), set by CMS [25], are shown as thin lines for certain  $M_{Z'}$  values. For specific sets of charges (labelled by  $E_6$ , GSM and LR, and described in [7]) parametrized by a mixing angle, the lower mass limit is given by the intersection of thick and thin lines. The black dots with smaller labels represent particular models.

It is common to present results of  $Z'$  searches as limits on the cross section versus  $M_{Z'}$  (see for example Fig. 1). An alternative is to plot exclusion curves for fixed  $M_{Z'}$  values in the  $c_u^\ell - c_d^\ell$  planes, allowing a simple derivation of the mass limit within any  $Z'$  model. The CMS upper limits in the  $c_u^\ell - c_d^\ell$  plane ( $\ell = e$  or  $\mu$ ) for different  $M_{Z'}$  are shown in Fig. 2 (for Tevatron limits, see [18,7]).

The discovery of a dilepton resonance at the LHC would determine the  $Z'$  mass and width. A measurement of the total cross section would define a band in the  $c_u^\ell - c_d^\ell$  plane. Angular distributions can be used to measure several combinations of  $Z'$  parameters (an example of how angular distributions improve the Tevatron sensitivity is given in [19]). Even though the original quark direction in a  $pp$  collider is unknown, the leptonic forward-backward asymmetry  $A_{FB}^\ell$  can be extracted from the kinematics of the dilepton system, and is sensitive to parity-violating couplings. A fit to the  $Z'$  rapidity distribution can distinguish between the couplings to up and down quarks. These measurements, combined with off-peak observables, have the potential to differentiate among various  $Z'$  models [20]. With  $100 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 14 \text{ TeV}$ , the spin of the  $Z'$  boson may be determined for  $M_{Z'} \leq 3 \text{ TeV}$  [21], and the expected sensitivity extends to  $M_{Z'} \sim 4 - 5 \text{ TeV}$  for many models [22].

Searches for  $Z'$  decays to  $e^+e^-$  and  $\mu^+\mu^-$  by the ATLAS and CMS collaborations set 95% C.L. lower limits on the  $Z'$  mass in the range between 2.8 and 3.4 TeV, depending on the specific model [23,24]. Lower mass limits for the flavor-violating leptonic final states have also been reported by ATLAS and CMS [26]; the limits obtained at 13 TeV in the  $e^\pm\mu^\mp$  channel are similar to those in the lepton-conserving channels above. In the case of final states with taus, lower limits obtained at 8 TeV are  $\approx 2.0 \text{ TeV}$  for the  $\tau^+\tau^-$  [27] decay and  $\approx 2.2 \text{ TeV}$  for the flavor-violating decays  $e^\pm\tau^\mp$  and  $\mu^\pm\tau^\mp$ .

Final states with higher background,  $t\bar{t}$ ,  $b\bar{b}$  and  $jj$ , are also important as they probe various combinations of  $Z'$  couplings to quarks. In the  $t\bar{t}$  channel, the 8 TeV data [28] sets lower mass limits in the 2–2.5 TeV range in a model where  $Z'$  couples only to the quarks of the first and third generations [29]. In the  $jj$  channel, the 13 TeV data [30] has been used to set limits on the production cross section of  $Z'$  bosons of masses larger than 1.5 TeV, where the trigger efficiency has reached its asymptotic value. For a comparison of earlier dijet resonance searches, see [31]. In the  $b\bar{b}$  channel [32], the  $b$  tagging leads to a reduction in both the background and the signal, so it may prove useful only if the  $Z' \rightarrow b\bar{b}$  branching fraction is large.

$Z'$  decays to  $Zh^0$  with  $Z \rightarrow \ell^+\ell^-$  or  $\nu\bar{\nu}$  and  $h^0 \rightarrow b\bar{b}$  have been studied by ATLAS [33] using 13 TeV data. The lower mass limit obtained in the context of the Heavy Vector Triplet model weakly-coupled scenario A [34] is 1.48 TeV. The  $Zh^0$  channel with the  $Z$  decaying hadronically and the Higgs boson decaying either hadronically or into  $\tau^+\tau^-$  has been studied by CMS [35] using 8 TeV data.

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The  $pp \rightarrow Z'X \rightarrow W^+W^-X$  process has also been searched for at the LHC. The channel where the  $Z'$  boson is produced through its couplings to quarks, and the  $W$  bosons decay hadronically, has been explored using boosted techniques to analyze the 13 TeV data [36]. The  $Z'$  boson may also be produced through its couplings to  $W$  bosons [37].

At the Tevatron, the CDF and DØ collaborations have searched for  $Z'$  bosons in the  $e^+e^-$  [38],  $\mu^+\mu^-$  [39],  $e^\pm\mu^\mp$  [40],  $\tau^+\tau^-$  [41],  $t\bar{t}$  [42],  $jj$  [43] and  $W^+W^-$  [44] final states. Although these limits have been often superseded by the LHC results, the Tevatron limits on certain  $Z'$  couplings (most notably, those arising from  $jj$  resonance searches [31]) remain competitive for  $M_{Z'}$  below about 0.5 TeV.

**Table 2:** Lower mass limits (in GeV) at 95% C.L. on various  $Z'$  bosons. The electroweak results [3] from low energy and  $W$  and  $Z$  boson data are for Higgs sectors consisting of doublets and singlets only ( $\rho_0 = 1$ ). The gauge coupling is fixed by an  $SO(10)$  unification condition for  $U(1)_\chi$ ,  $U(1)_\psi$  and  $U(1)_\eta$ . The secluded  $Z_S$  emerges in a supersymmetric model [46], and  $Z_{SSM}$  is the sequential  $Z'$  (same coupling as the SM  $Z$  boson). The last three columns show the limits from dilepton resonance searches at the LHC [23,24] and the Tevatron [39,38], and from  $e^+e^- \rightarrow f\bar{f}$  measurements at LEP-II [16].

$Z'$	electroweak	ATLAS/CMS	CDF/DØ	LEP-II
$Z_\chi$	1141	3080	930	785
$Z_\psi$	147	2790	917	500
$Z_\eta$	427	2850	938	500
$Z_S$	1257	3030	858	—
$Z_{SSM}$	1403	3400	1071	1760

**Low-energy constraints.**  $Z'$  boson properties are also constrained by a variety of low-energy experiments [45]. Polarized electron-nucleon scattering and atomic parity violation are sensitive to electron-quark contact interactions, which get contributions from  $Z'$  exchange that can be expressed in terms of the couplings introduced in Eq. (1) and  $M_{Z'}$ . Further corrections to the electron-quark contact interactions are induced in the presence of  $\tilde{Z} - Z'$  mixing because of the shifts in the  $Z$  couplings to quarks and leptons [2]. Deep-inelastic neutrino-nucleon scattering is similarly affected by  $Z'$  bosons. Other low-energy observables are discussed in [3]. In some models, the lower limits on  $M_{Z'}$  set by low energy data are above 1 TeV, as shown in Table 2 (for more general models, see [1,5,47]). The mass bounds from direct searches at the LHC [23,24] exceed the electroweak constraints by a factor of two or more for the models considered there. While the electroweak constraints can be slightly improved by fixing the Higgs mass to the value

measured at the LHC, and the collider bounds are moderately weakened if there are open exotic decay channels [48], this conclusion will not change.

Although the LHC data are most constraining for many  $Z'$  models, one should be careful in assessing the relative reach of various experiments given the freedom in  $Z'$  couplings. For example, a  $Z'$  coupled to  $B - yL_\mu + (y - 3)L_\tau$  has implications for the muon  $g - 2$ , neutrino oscillations or  $\tau$  decays, and would be hard to see in processes involving first-generation fermions. Moreover, the combination of LHC searches and low-energy measurements could allow a precise determination of the  $Z'$  parameters [49].

## References

- For reviews, see P. Langacker, Rev. Mod. Phys. **81**, 1199 (2009); A. Leike, Phys. Rept. **317**, 143 (1999); J. Hewett and T. Rizzo, Phys. Rept. **183**, 193 (1989).
- K.S. Babu *et al.*, Phys. Rev. **D57**, 6788 (1998); B. Holdom, Phys. Lett. **B259**, 329 (1991).
- J. Erler *et al.*, JHEP **0908**, 017 (2009).
- B. A. Dobrescu and P. J. Fox, arXiv:1511.02148.
- M.S. Carena *et al.*, Phys. Rev. **D70**, 093009 (2004).
- See, e.g., F. Del Aguila *et al.*, Phys. Rev. **D52**, 37 (1995).
- E. Accomando *et al.*, Phys. Rev. **D83**, 075012 (2011).
- M.-C. Chen *et al.*, Phys. Rev. **D75**, 055009 (2007).
- P. Langacker and M. Plumacher, Phys. Rev. **D62**, 013006 (2000); R.S. Chivukula and E.H. Simmons, Phys. Rev. **D66**, 015006 (2002).
- A. J. Buras *et al.*, JHEP **1302**, 116 (2013).
- L.E. Ibanez and G.G. Ross, Phys. Lett. **B332**, 100 (1994).
- See the Section on “ $W'$  searches” in this Review.
- J. Parsons and A. Pomarol, “Extra dimensions” in this Review.
- R.S. Chivukula *et al.*, “Dynamical electroweak symmetry breaking” in this Review.
- M. Bando *et al.*, Phys. Rept. **164**, 217 (1988).
- S. Schael *et al.* [ALEPH, DELPHI, L3, OPAL and LEP Electroweak Collaborations], Phys. Rept. **532**, 119 (2013).
- E. Accomando *et al.*, JHEP **1310**, 153 (2013).
- A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **95**, 252001 (2005).
- A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **96**, 211801 (2006).
- F. Petriello and S. Quackenbush, Phys. Rev. **D77**, 115004 (2008).
- P. Osland *et al.*, Phys. Rev. **D79**, 115021 (2009).
- G.L. Bayatian *et al.* [CMS Collab.], J. Phys. **G34**, 995 (2007).
- ATLAS Collab., note CONF-2015-070, Dec. 2015.
- CMS Collab., note PAS-EXO-15-005, Dec. 2015.
- V. Khachatryan *et al.* [CMS Collab.], JHEP **1504**, 025 (2015).
- ATLAS Collab., note CONF-2015-072, Dec. 2015; G. Aad *et al.* [ATLAS Collab.], Phys. Rev. Lett. **115**, 031801 (2015); CMS Collab., note PAS-EXO-13-002, Mar. 2015.
- G. Aad *et al.* [ATLAS Collab.], JHEP **1507**, 157 (2015); CMS Collab., note PAS-EXO-12-046, Jan. 2015.

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

28. G. Aad *et al.* [ATLAS Collab.], JHEP **1508**, 148 (2015); V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. D **93**, 012001 (2016).
29. R. M. Harris, C. T. Hill and S. J. Parke, hep-ph/9911288..
30. G. Aad *et al.* [ATLAS Collab.], arXiv:1512.01530; Phys. Rev. D **91**, 052007 (2015); V. Khachatryan *et al.* [CMS Collab.], arXiv:1512.01224; Phys. Rev. D **91**, 052009 (2015); CMS Collab., note PAS-EXO-14-005, Oct. 2015.
31. B.A. Dobrescu and F. Yu, Phys. Rev. D **88**, 035021 (2013).
32. CMS Collab., note PAS-EXO-12-023, Apr. 2013.
33. ATLAS Collab., note CONF-2015-074, Dec. 2015.
34. D. Pappadopulo *et al.*, JHEP **1409**, 060 (2014).
35. V. Khachatryan *et al.* [CMS Collab.], arXiv:1506.01443 and Phys. Lett. B **748**, 255 (2015).
36. ATLAS Collab., note CONF-2015-073, Dec. 2015.
37. H. J. He *et al.*, Phys. Rev. D **78**, 031701 (2008).
38. V. M. Abazov *et al.* [D0 Collab.], Phys. Lett. B **695**, 88 (2011); T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. Lett. **102**, 031801 (2009).
39. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. Lett. **106**, 121801 (2011).
40. A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **96**, 211802 (2006); V. M. Abazov *et al.* [D0 Collab.], Phys. Rev. Lett. **105**, 191802 (2010).
41. D. Acosta *et al.* [CDF Collab.], Phys. Rev. Lett. **95**, 131801 (2005).
42. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. D **84**, 072004 (2011); V. M. Abazov *et al.* [D0 Collab.], Phys. Rev. D **85**, 051101 (2012).
43. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. D **79**, 112002 (2009).
44. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. Lett. **104**, 241801 (2010).
45. See, e.g., V.D. Barger *et al.*, Phys. Rev. D **57**, 391 (1998); J. Erler and M.J. Ramsey-Musolf, Prog. Part. Nucl. Phys. **54**, 351 (2005).
46. J. Erler *et al.*, Phys. Rev. D **66**, 015002 (2002); G. Cleaver *et al.*, Phys. Rev. D **59**, 055005 (1999).
47. E. Rojas and J. Erler, JHEP **1510**, 063 (2015).
48. J. Kang and P. Langacker, Phys. Rev. D **71**, 035014 (2005); C.-F. Chang *et al.*, JHEP **1109**, 058 (2011).
49. Y. Li *et al.*, Phys. Rev. D **80**, 055018 (2009).

### MASS LIMITS for $Z'$ (Heavy Neutral Vector Boson Other Than $Z$ )

#### Limits for $Z'_{SM}$

$Z'_{SM}$  is assumed to have couplings with quarks and leptons which are identical to those of  $Z$ , and decays only to known fermions. The most recent preliminary results can be found in the "Z'-boson searches" review above.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>2020	95	1 AAD	15AMATLS	$pp; Z'_{SM} \rightarrow \tau^+ \tau^-$
<b>&gt;2900</b>	95	2 KHACHATRYAN...15AE	CMS	$pp; Z'_{SM} \rightarrow e^+ e^-, \mu^+ \mu^-$
none 1200–1700	95	3 KHACHATRYAN...15V	CMS	$pp; Z'_{SM} \rightarrow q\bar{q}$
<b>&gt;2900</b>	95	4 AAD	14V ATLS	$pp; Z'_{SM} \rightarrow e^+ e^-, \mu^+ \mu^-$
>1470	95	5 CHATRCHYAN13A	CMS	$pp; Z'_{SM} \rightarrow q\bar{q}$
>1400	95	6 CHATRCHYAN120	CMS	$pp; Z'_{SM} \rightarrow \tau^+ \tau^-$
<b>&gt;1500</b>	95	7 CHEUNG	01B RVUE	Electroweak

• • • We do not use the following data for averages, fits, limits, etc. • • •

>1400	95	8 AAD	13s ATLS	$pp; Z'_{SM} \rightarrow \tau^+ \tau^-$
>2590	95	9 CHATRCHYAN13AF	CMS	$pp; Z'_{SM} \rightarrow e^+ e^-, \mu^+ \mu^-$
>2220	95	10 AAD	12cc ATLS	$pp; Z'_{SM} \rightarrow e^+ e^-, \mu^+ \mu^-$
>1071	95	11 AALTONEN	11i CDF	$p\bar{p}; Z'_{SM} \rightarrow \mu^+ \mu^-$
>1023	95	12 ABZOV	11A D0	$p\bar{p}; Z'_{SM} \rightarrow e^+ e^-$
none 247–544	95	13 AALTONEN	10N CDF	$Z' \rightarrow WW$
none 320–740	95	14 AALTONEN	09AC CDF	$Z' \rightarrow q\bar{q}$
> 963	95	12 AALTONEN	09T CDF	$p\bar{p}; Z'_{SM} \rightarrow e^+ e^-$
>1403	95	15 ERLER	09 RVUE	Electroweak
>1305	95	16 ABDALLAH	06C DLPH	$e^+ e^-$
> 399	95	17 ACOSTA	05R CDF	$p\bar{p}; Z'_{SM} \rightarrow \tau^+ \tau^-$
none 400–640	95	ABAZOV	04C D0	$p\bar{p}; Z'_{SM} \rightarrow q\bar{q}$
>1018	95	18 ABBIENDI	04G OPAL	$e^+ e^-$
> 670	95	19 ABZOV	01B D0	$p\bar{p}; Z'_{SM} \rightarrow e^+ e^-$
> 710	95	20 ABREU	00S DLPH	$e^+ e^-$
> 898	95	21 BARATE	00i ALEP	$e^+ e^-$
> 809	95	22 ERLER	99 RVUE	Electroweak
> 690	95	23 ABE	97S CDF	$p\bar{p}; Z'_{SM} \rightarrow e^+ e^-, \mu^+ \mu^-$
> 398	95	24 VILAIN	94B CHM2	$\nu_\mu e \rightarrow \nu_\mu e + \bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$
> 237	90	25 ALITTI	93 UA2	$p\bar{p}; Z'_{SM} \rightarrow q\bar{q}$
none 260–600	95	26 RIZZO	93 RVUE	$p\bar{p}; Z'_{SM} \rightarrow q\bar{q}$
> 426	90	27 ABE	90F VNS	$e^+ e^-$

- 1 AAD 15AM search for resonances decaying to  $\tau^+ \tau^-$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- 2 KHACHATRYAN 15AE search for resonances decaying to  $e^+ e^-, \mu^+ \mu^-$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- 3 KHACHATRYAN 15V search for resonances decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- 4 AAD 14v search for resonances decaying to  $e^+ e^-, \mu^+ \mu^-$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- 5 CHATRCHYAN 13A use  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 6 CHATRCHYAN 120 search for resonances decaying to  $\tau^+ \tau^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 7 CHEUNG 01B limit is derived from bounds on contact interactions in a global electroweak analysis.
- 8 AAD 13s search for resonances decaying to  $\tau^+ \tau^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 9 CHATRCHYAN 13AF search for resonances decaying to  $e^+ e^-, \mu^+ \mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV and 8 TeV.
- 10 AAD 12cc search for resonances decaying to  $e^+ e^-, \mu^+ \mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 11 AALTONEN 11i search for resonances decaying to  $\mu^+ \mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 12 ABZOV 11A, AALTONEN 09T, AALTONEN 07H, and ABULENCIA 06L search for resonances decaying to  $e^+ e^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 13 The quoted limit assumes  $g_{WWZ'}/g_{WWZ} = (M_W/M_{Z'})^2$ . See their Fig. 4 for limits in mass-coupling plane.
- 14 AALTONEN 09AC search for new particle decaying to dijets.
- 15 ERLER 09 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0026 < \theta < 0.0006$ .
- 16 ABDALLAH 06C use data  $\sqrt{s} = 130$ –207 GeV.
- 17 ACOSTA 05R search for resonances decaying to tau lepton pairs in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 18 ABBIENDI 04G give 95% CL limit on  $Z$ - $Z'$  mixing  $-0.00422 < \theta < 0.00091$ .  $\sqrt{s} = 91$  to 207 GeV.
- 19 ABZOV 01B search for resonances in  $p\bar{p} \rightarrow e^+ e^-$  at  $\sqrt{s} = 1.8$  TeV. They find  $\sigma \cdot B(Z' \rightarrow e\bar{e}) < 0.06$  pb for  $M_{Z'} > 500$  GeV.
- 20 ABREU 00S uses LEP data at  $\sqrt{s} = 90$  to 189 GeV.
- 21 BARATE 00i search for deviations in cross section and asymmetries in  $e^+ e^- \rightarrow$  fermions at  $\sqrt{s} = 90$  to 183 GeV. Assume  $\theta = 0$ . Bounds in the mass-mixing plane are shown in their Figure 18.
- 22 ERLER 99 give 90%CL limit on the  $Z$ - $Z'$  mixing  $-0.0041 < \theta < 0.0003$ .  $\rho_0 = 1$  is assumed.
- 23 ABE 97S find  $\sigma(Z') \times B(e^+ e^-, \mu^+ \mu^-) < 40$  fb for  $m_{Z'} > 600$  GeV at  $\sqrt{s} = 1.8$  TeV.
- 24 VILAIN 94B assume  $m_t = 150$  GeV.
- 25 ALITTI 93 search for resonances in the two-jet invariant mass. The limit assumes  $B(Z' \rightarrow q\bar{q}) = 0.7$ . See their Fig. 5 for limits in the  $m_{Z'} - B(q\bar{q})$  plane.
- 26 RIZZO 93 analyses CDF limit on possible two-jet resonances.
- 27 ABE 90F use data for  $R, R_{\ell\ell}$ , and  $A_{\ell\ell}$ . They fix  $m_W = 80.49 \pm 0.43 \pm 0.24$  GeV and  $m_Z = 91.13 \pm 0.03$  GeV.

#### Limits for $Z_{LR}$

$Z_{LR}$  is the extra neutral boson in left-right symmetric models.  $g_L = g_R$  is assumed unless noted. Values in parentheses assume stronger constraint on the Higgs sector, usually motivated by specific left-right symmetric models (see the Note on the  $W'$ ). Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino. Direct search bounds assume decays to Standard Model fermions only, unless noted.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;1162</b>	95	1 DEL-AGUILA	10 RVUE	Electroweak
<b>&gt; 630</b>	95	2 ABE	97S CDF	$p\bar{p}; Z'_{LR} \rightarrow e^+ e^-, \mu^+ \mu^-$

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 998	95	<sup>3</sup> ERLER	09	RVUE	Electroweak
> 600	95	SCHAEEL	07A	ALEP	$e^+e^-$
> 455	95	<sup>4</sup> ABDALLAH	06c	DLPH	$e^+e^-$
> 518	95	<sup>5</sup> ABBIENDI	04c	OPAL	$e^+e^-$
> 860	95	<sup>6</sup> CHEUNG	01B	RVUE	Electroweak
> 380	95	<sup>7</sup> ABREU	00S	DLPH	$e^+e^-$
> 436	95	<sup>8</sup> BARATE	00i	ALEP	Repl. by SCHAEEL 07A
> 550	95	<sup>9</sup> CHAY	00	RVUE	Electroweak
		<sup>10</sup> ERLER	00	RVUE	Cs
		<sup>11</sup> CASALBUONI	99	RVUE	Cs
(> 1205)	90	<sup>12</sup> CZAKON	99	RVUE	Electroweak
> 564	95	<sup>13</sup> ERLER	99	RVUE	Electroweak
(> 1673)	95	<sup>14</sup> ERLER	99	RVUE	Electroweak
(> 1700)	68	<sup>15</sup> BARENBOIM	98	RVUE	Electroweak
> 244	95	<sup>16</sup> CONRAD	98	RVUE	$\nu_\mu N$ scattering
> 253	95	<sup>17</sup> VILAIN	94B	CHM2	$\nu_\mu e \rightarrow \nu_\mu e$ and $\bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$
none 200–600	95	<sup>18</sup> RIZZO	93	RVUE	$p\bar{p}; Z_{LR} \rightarrow q\bar{q}$
[> 2000]		<sup>19</sup> WALKER	91	COSM	Nucleosynthesis; light $\nu_R$
none 200–500		<sup>20</sup> GRIFOLS	90	ASTR	SN 1987A; light $\nu_R$
none 350–2400		<sup>21</sup> BARBIERI	89B	ASTR	SN 1987A; light $\nu_R$

<sup>1</sup> DEL-AGUILA 10 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0012 < \theta < 0.0004$ .

<sup>2</sup> ABE 97s find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40 \text{ fb}$  for  $m_{Z'} > 600 \text{ GeV}$  at  $\sqrt{s} = 1.8 \text{ TeV}$ .

<sup>3</sup> ERLER 09 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0013 < \theta < 0.0006$ .

<sup>4</sup> ABDALLAH 06c give 95% CL limit  $|\theta| < 0.0028$ . See their Fig. 14 for limit contours in the mass-mixing plane.

<sup>5</sup> ABBIENDI 04c give 95% CL limit on  $Z$ - $Z'$  mixing  $-0.00098 < \theta < 0.00190$ . See their Fig. 20 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 91$  to 207 GeV.

<sup>6</sup> CHEUNG 01B limit is derived from bounds on contact interactions in a global electroweak analysis.

<sup>7</sup> ABREU 00s give 95% CL limit on  $Z$ - $Z'$  mixing  $|\theta| < 0.0018$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 90$  to 189 GeV.

<sup>8</sup> BARATE 00i search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s} = 90$  to 183 GeV. Assume  $\theta = 0$ . Bounds in the mass-mixing plane are shown in their Figure 18.

<sup>9</sup> CHAY 00 also find  $-0.0003 < \theta < 0.0019$ . For  $g_R$  free,  $m_{Z'} > 430 \text{ GeV}$ .

<sup>10</sup> ERLER 00 discuss the possibility that a discrepancy between the observed and predicted values of  $Q_W(\text{Cs})$  is due to the exchange of  $Z'$ . The data are better described in a certain class of the  $Z'$  models including  $Z_{LR}$  and  $Z_\chi$ .

<sup>11</sup> CASALBUONI 99 discuss the discrepancy between the observed and predicted values of  $Q_W(\text{Cs})$ . It is shown that the data are better described in a class of models including the  $Z_{LR}$  model.

<sup>12</sup> CZAKON 99 perform a simultaneous fit to charged and neutral sectors. Assumes manifest left-right symmetric model. Finds  $|\theta| < 0.0042$ .

<sup>13</sup> ERLER 99 give 90% CL limit on the  $Z$ - $Z'$  mixing  $-0.0009 < \theta < 0.0017$ .

<sup>14</sup> ERLER 99 assumes 2 Higgs doublets, transforming as 10 of  $\text{SO}(10)$ , embedded in  $E_6$ .

<sup>15</sup> BARENBOIM 98 also gives 68% CL limits on the  $Z$ - $Z'$  mixing  $-0.0005 < \theta < 0.0033$ . Assumes Higgs sector of minimal left-right model.

<sup>16</sup> CONRAD 98 limit is from measurements at CCFR, assuming no  $Z$ - $Z'$  mixing.

<sup>17</sup> VILAIN 94B assume  $m_t = 150 \text{ GeV}$  and  $\theta = 0$ . See Fig. 2 for limit contours in the mass-mixing plane.

<sup>18</sup> RIZZO 93 analyses CDF limit on possible two-jet resonances.

<sup>19</sup> GRIFOLS 90 limit holds for  $m_{\nu_R} \lesssim 1 \text{ MeV}$ . A specific Higgs sector is assumed. See also GRIFOLS 90D, RIZZO 91.

<sup>20</sup> BARBIERI 89B limit holds for  $m_{\nu_R} \lesssim 10 \text{ MeV}$ . Bounds depend on assumed supernova core temperature.

### Limits for $Z_\chi$

$Z_\chi$  is the extra neutral boson in  $\text{SO}(10) \rightarrow \text{SU}(5) \times \text{U}(1)_\chi$ .  $g_\chi = e/\cos\theta_W$  is assumed unless otherwise stated. We list limits with the assumption  $\rho = 1$  but with no further constraints on the Higgs sector. Values in parentheses assume stronger constraint on the Higgs sector motivated by superstring models. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>2620	95	<sup>1</sup> AAD	14v ATLS	$p\bar{p}, Z'_\chi \rightarrow e^+e^-, \mu^+\mu^-$
>1141	95	<sup>2</sup> ERLER	09 RVUE	Electroweak
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>1970	95	<sup>3</sup> AAD	12cc ATLS	$p\bar{p}, Z'_\chi \rightarrow e^+e^-, \mu^+\mu^-$
> 930	95	<sup>4</sup> AALTONEN	11i CDF	$p\bar{p}; Z'_\chi \rightarrow \mu^+\mu^-$
> 903	95	<sup>5</sup> ABAZOV	11A D0	$p\bar{p}, Z'_\chi \rightarrow e^+e^-$
>1022	95	<sup>6</sup> DEL-AGUILA	10 RVUE	Electroweak
> 862	95	<sup>5</sup> AALTONEN	09t CDF	$p\bar{p}, Z'_\chi \rightarrow e^+e^-$
> 892	95	<sup>7</sup> AALTONEN	09v CDF	Repl. by AALTONEN 11i
> 822	95	<sup>5</sup> AALTONEN	07H CDF	Repl. by AALTONEN 09t
> 680	95	SCHAEEL	07A ALEP	$e^+e^-$
> 545	95	<sup>8</sup> ABDALLAH	06c DLPH	$e^+e^-$
> 740	95	<sup>5</sup> ABULENCIA	06L CDF	Repl. by AALTONEN 07H
> 690	95	<sup>9</sup> ABULENCIA	05A CDF	$p\bar{p}; Z'_\chi \rightarrow e^+e^-, \mu^+\mu^-$
> 781	95	<sup>10</sup> ABBIENDI	04G OPAL	$e^+e^-$
>2100		<sup>11</sup> BARGER	03B COSM	Nucleosynthesis; light $\nu_R$
> 680	95	<sup>12</sup> CHEUNG	01B RVUE	Electroweak
> 440	95	<sup>13</sup> ABREU	00S DLPH	$e^+e^-$

> 533	95	<sup>14</sup> BARATE	00i ALEP	Repl. by SCHAEEL 07A
> 554	95	<sup>15</sup> CHO	00 RVUE	Electroweak
		<sup>16</sup> ERLER	00 RVUE	Cs
		<sup>17</sup> ROSNER	00 RVUE	Cs
> 545	95	<sup>18</sup> ERLER	99 RVUE	Electroweak
(> 1368)	95	<sup>19</sup> ERLER	99 RVUE	Electroweak
> 215	95	<sup>20</sup> CONRAD	98 RVUE	$\nu_\mu N$ scattering
> 595	95	<sup>21</sup> ABE	97s CDF	$p\bar{p}; Z'_\chi \rightarrow e^+e^-, \mu^+\mu^-$
> 190	95	<sup>22</sup> ARIMA	97 VNS	Bhabha scattering
> 262	95	<sup>23</sup> VILAIN	94B CHM2	$\nu_\mu e \rightarrow \nu_\mu e; \bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$
[>1470]		<sup>24</sup> FARAGGI	91 COSM	Nucleosynthesis; light $\nu_R$
> 231	90	<sup>25</sup> ABE	90f VNS	$e^+e^-$
[> 1140]		<sup>26</sup> GONZALEZ-G.	90D COSM	Nucleosynthesis; light $\nu_R$
[> 2100]		<sup>27</sup> GRIFOLS	90 ASTR	SN 1987A; light $\nu_R$

<sup>1</sup> AAD 14v search for resonances decaying to  $e^+e^-, \mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$ .

<sup>2</sup> ERLER 09 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0016 < \theta < 0.0006$ .

<sup>3</sup> AAD 12cc search for resonances decaying to  $e^+e^-, \mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$ .

<sup>4</sup> AALTONEN 11i search for resonances decaying to  $\mu^+\mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .

<sup>5</sup> ABAZOV 11A, AALTONEN 09t, AALTONEN 07H, and ABULENCIA 06L search for resonances decaying to  $e^+e^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .

<sup>6</sup> DEL-AGUILA 10 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0011 < \theta < 0.0007$ .

<sup>7</sup> AALTONEN 09v search for resonances decaying to  $\mu^+\mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .

<sup>8</sup> ABDALLAH 06c give 95% CL limit  $|\theta| < 0.0031$ . See their Fig. 14 for limit contours in the mass-mixing plane.

<sup>9</sup> ABULENCIA 05A search for resonances decaying to electron or muon pairs in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .

<sup>10</sup> ABBIENDI 04c give 95% CL limit on  $Z$ - $Z'$  mixing  $-0.00099 < \theta < 0.00194$ . See their Fig. 20 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 91$  to 207 GeV.

<sup>11</sup> BARGER 03B limit is from the nucleosynthesis bound on the effective number of light neutrino  $\delta N_\nu < 1$ . The quark-hadron transition temperature  $T_C = 150 \text{ MeV}$  is assumed. The limit with  $T_C = 400 \text{ MeV}$  is  $> 4300 \text{ GeV}$ .

<sup>12</sup> CHEUNG 01B limit is derived from bounds on contact interactions in a global electroweak analysis.

<sup>13</sup> ABREU 00s give 95% CL limit on  $Z$ - $Z'$  mixing  $|\theta| < 0.0017$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 90$  to 189 GeV.

<sup>14</sup> BARATE 00i search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s} = 90$  to 183 GeV. Assume  $\theta = 0$ . Bounds in the mass-mixing plane are shown in their Figure 18.

<sup>15</sup> CHO 00 use various electroweak data to constrain  $Z'$  models assuming  $m_H = 100 \text{ GeV}$ . See Fig. 3 for limits in the mass-mixing plane.

<sup>16</sup> ERLER 00 discuss the possibility that a discrepancy between the observed and predicted values of  $Q_W(\text{Cs})$  is due to the exchange of  $Z'$ . The data are better described in a certain class of the  $Z'$  models including  $Z_{LR}$  and  $Z_\chi$ .

<sup>17</sup> ROSNER 00 discusses the possibility that a discrepancy between the observed and predicted values of  $Q_W(\text{Cs})$  is due to the exchange of  $Z'$ . The data are better described in a certain class of the  $Z'$  models including  $Z_\chi$ .

<sup>18</sup> ERLER 99 give 90% CL limit on the  $Z$ - $Z'$  mixing  $-0.0020 < \theta < 0.0015$ .

<sup>19</sup> ERLER 99 assumes 2 Higgs doublets, transforming as 10 of  $\text{SO}(10)$ , embedded in  $E_6$ .

<sup>20</sup> CONRAD 98 limit is from measurements at CCFR, assuming no  $Z$ - $Z'$  mixing.

<sup>21</sup> ABE 97s find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40 \text{ fb}$  for  $m_{Z'} > 600 \text{ GeV}$  at  $\sqrt{s} = 1.8 \text{ TeV}$ .

<sup>22</sup>  $Z$ - $Z'$  mixing is assumed to be zero.  $\sqrt{s} = 57.77 \text{ GeV}$ .

<sup>23</sup> VILAIN 94B assume  $m_t = 150 \text{ GeV}$  and  $\theta = 0$ . See Fig. 2 for limit contours in the mass-mixing plane.

<sup>24</sup> FARAGGI 91 limit assumes the nucleosynthesis bound on the effective number of neutrinos  $\Delta N_\nu < 0.5$  and is valid for  $m_{\nu_R} < 1 \text{ MeV}$ .

<sup>25</sup> ABE 90f use data for  $R, R_{\ell\ell}$ , and  $A_{\ell\ell}$ . ABE 90f fix  $m_W = 80.49 \pm 0.43 \pm 0.24 \text{ GeV}$  and  $m_Z = 91.13 \pm 0.03 \text{ GeV}$ .

<sup>26</sup> Assumes the nucleosynthesis bound on the effective number of light neutrinos ( $\delta N_\nu < 1$ ) and that  $\nu_R$  is light ( $\lesssim 1 \text{ MeV}$ ).

<sup>27</sup> GRIFOLS 90 limit holds for  $m_{\nu_R} \lesssim 1 \text{ MeV}$ . See also GRIFOLS 90D, RIZZO 91.

### Limits for $Z_\psi$

$Z_\psi$  is the extra neutral boson in  $E_6 \rightarrow \text{SO}(10) \times \text{U}(1)_\psi$ .  $g_\psi = e/\cos\theta_W$  is assumed unless otherwise stated. We list limits with the assumption  $\rho = 1$  but with no further constraints on the Higgs sector. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>2570	95	<sup>1</sup> KHACHATRYAN	15AE CMS	$p\bar{p}; Z'_\psi \rightarrow e^+e^-, \mu^+\mu^-$
>2510	95	<sup>2</sup> AAD	14v ATLS	$p\bar{p}, Z'_\psi \rightarrow e^+e^-, \mu^+\mu^-$
>1100	95	<sup>3</sup> CHATRCHYAN	12a CMS	$p\bar{p}, Z'_\psi \rightarrow \tau^+\tau^-$
> 476	95	<sup>4</sup> DEL-AGUILA	10 RVUE	Electroweak
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>2260	95	<sup>5</sup> CHATRCHYAN	13AF CMS	$p\bar{p}, Z'_\psi \rightarrow e^+e^-, \mu^+\mu^-$
>1790	95	<sup>6</sup> AAD	12cc ATLS	$p\bar{p}, Z'_\psi \rightarrow e^+e^-, \mu^+\mu^-$
>2000	95	<sup>7</sup> CHATRCHYAN	12m CMS	Repl. by CHATRCHYAN 13AF
> 917	95	<sup>8</sup> AALTONEN	11i CDF	$p\bar{p}; Z'_\psi \rightarrow \mu^+\mu^-$
> 891	95	<sup>9</sup> ABAZOV	11A D0	$p\bar{p}, Z'_\psi \rightarrow e^+e^-$
> 851	95	<sup>9</sup> AALTONEN	09t CDF	$p\bar{p}, Z'_\psi \rightarrow e^+e^-$
> 878	95	<sup>10</sup> AALTONEN	09v CDF	Repl. by AALTONEN 11i
> 147	95	<sup>11</sup> ERLER	09 RVUE	Electroweak

See key on page 601

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

> 822	95	9	AALTONEN	07H	CDF	Repl. by AALTONEN 09T
> 410	95		SCHAEEL	07A	ALEP	$e^+e^-$
> 475	95	12	ABDALLAH	06C	DLPH	$e^+e^-$
> 725		9	ABULENCIA	06L	CDF	Repl. by AALTONEN 07H
> 675	95	13	ABULENCIA	05A	CDF	Repl. by AALTONEN 11I and AALTONEN 09T
> 366	95	14	ABBIENDI	04G	OPAL	$e^+e^-$
> 600		15	BARGER	03B	COSM	Nucleosynthesis; light $\nu_R$
> 350	95	16	ABREU	00S	DLPH	$e^+e^-$
> 294	95	17	BARATE	00I	ALEP	Repl. by SCHAEEL 07A
> 137	95	18	CHO	00	RVUE	Electroweak
> 146	95	19	ERLER	99	RVUE	Electroweak
> 54	95	20	CONRAD	98	RVUE	$\nu_\mu N$ scattering
> 590	95	21	ABE	97S	CDF	$p\bar{p}; Z'_\psi \rightarrow e^+e^-, \mu^+\mu^-$
> 135	95	22	VILAIN	94B	CHM2	$\nu_\mu e \rightarrow \nu_\mu e; \bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$
> 105	90	23	ABE	90F	VNS	$e^+e^-$
[> 160]		24	GONZALEZ-G.	90D	COSM	Nucleosynthesis; light $\nu_R$
[> 2000]		25	GRIFOLS	90D	ASTR	SN 1987A; light $\nu_R$

- 1 KHACHATRYAN 15AE search for resonances decaying to  $e^+e^-$ ,  $\mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- 2 AAD 14V search for resonances decaying to  $e^+e^-$ ,  $\mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- 3 CHATRCHYAN 12o search for resonances decaying to  $\tau^+\tau^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 4 DEL-AGUILA 10 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0019 < \theta < 0.0007$ .
- 5 CHATRCHYAN 13AF search for resonances decaying to  $e^+e^-$ ,  $\mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV and 8 TeV.
- 6 AAD 12CC search for resonances decaying to  $e^+e^-$ ,  $\mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 7 CHATRCHYAN 12M search for resonances decaying to  $e^+e^-$  or  $\mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 8 AALTONEN 11I search for resonances decaying to  $\mu^+\mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 9 ABZOV 11A, AALTONEN 09T, AALTONEN 07H, and ABULENCIA 06L search for resonances decaying to  $e^+e^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 10 AALTONEN 09V search for resonances decaying to  $\mu^+\mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 11 ERLER 09 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0018 < \theta < 0.0009$ .
- 12 ABDALLAH 06C give 95% CL limit  $|\theta| < 0.0027$ . See their Fig. 14 for limit contours in the mass-mixing plane.
- 13 ABULENCIA 05A search for resonances decaying to electron or muon pairs in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 14 ABBIENDI 04G give 95% CL limit on  $Z$ - $Z'$  mixing  $-0.00129 < \theta < 0.00258$ . See their Fig. 20 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 91$  to 207 GeV.
- 15 BARGER 03B limit is from the nucleosynthesis bound on the effective number of light neutrino  $\delta N_\nu < 1$ . The quark-hadron transition temperature  $T_C = 150$  MeV is assumed. The limit with  $T_C = 400$  MeV is  $> 1100$  GeV.
- 16 ABREU 00S give 95% CL limit on  $Z$ - $Z'$  mixing  $|\theta| < 0.0018$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 90$  to 189 GeV.
- 17 BARATE 00I search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s} = 90$  to 183 GeV. Assume  $\theta = 0$ . Bounds in the mass-mixing plane are shown in their Figure 18.
- 18 CHO 00 use various electroweak data to constrain  $Z'$  models assuming  $m_H = 100$  GeV. See Fig. 3 for limits in the mass-mixing plane.
- 19 ERLER 99 give 90% CL limit on the  $Z$ - $Z'$  mixing  $-0.0013 < \theta < 0.0024$ .
- 20 CONRAD 98 limit is from measurements at CCFR, assuming no  $Z$ - $Z'$  mixing.
- 21 ABE 97S find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40$  fb for  $m_{Z'} > 600$  GeV at  $\sqrt{s} = 1.8$  TeV.
- 22 VILAIN 94B assume  $m_t = 150$  GeV and  $\theta = 0$ . See Fig. 2 for limit contours in the mass-mixing plane.
- 23 ABE 90F use data for  $R$ ,  $R_{\ell\ell}$ , and  $A_{\ell\ell}$ . ABE 90F fix  $m_W = 80.49 \pm 0.43 \pm 0.24$  GeV and  $m_Z = 91.13 \pm 0.03$  GeV.
- 24 Assumes the nucleosynthesis bound on the effective number of light neutrinos ( $\delta N_\nu < 1$ ) and that  $\nu_R$  is light ( $\lesssim 1$  MeV).
- 25 GRIFOLS 90D limit holds for  $m_{\nu_R} \lesssim 1$  MeV. See also RIZZO 91.

### Limits for $Z_\eta$

$Z_\eta$  is the extra neutral boson in  $E_6$  models, corresponding to  $Q_\eta = \sqrt{3/8} Q_\chi - \sqrt{5/8} Q_\psi$ .  $g_\eta = e/\cos\theta_W$  is assumed unless otherwise stated. We list limits with the assumption  $\rho = 1$  but with no further constraints on the Higgs sector. Values in parentheses assume stronger constraint on the Higgs sector motivated by superstring models. Values in brackets are from cosmological and astrophysical considerations and assume a light right-handed neutrino.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 1870	95	1 AAD	12CC ATLS	$pp, Z'_\eta \rightarrow e^+e^-, \mu^+\mu^-$
> 619	95	2 CHO	00 RVUE	Electroweak
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 938	95	3 AALTONEN	11I CDF	$p\bar{p}; Z'_\eta \rightarrow \mu^+\mu^-$
> 923	95	4 ABZOV	11A D0	$p\bar{p}; Z'_\eta \rightarrow e^+e^-$
> 488	95	5 DEL-AGUILA	10 RVUE	Electroweak
> 877	95	6 AALTONEN	09T CDF	$p\bar{p}; Z'_\eta \rightarrow e^+e^-$
> 904	95	7 ERLER	09 RVUE	Electroweak
> 427	95	8 AALTONEN	07H CDF	Repl. by AALTONEN 09T
> 891	95	9 SCHAEEL	07A ALEP	$e^+e^-$
> 350	95	10 ABDALLAH	06C DLPH	$e^+e^-$
> 360	95	11 ABULENCIA	06L CDF	Repl. by AALTONEN 07H
> 745	95	12 ABULENCIA	05A CDF	Repl. by AALTONEN 11I and AALTONEN 09T

> 515	95	10	ABBIENDI	04G	OPAL	$e^+e^-$
> 1600		11	BARGER	03B	COSM	Nucleosynthesis; light $\nu_R$
> 310	95	12	ABREU	00S	DLPH	$e^+e^-$
> 329	95	13	BARATE	00I	ALEP	Repl. by SCHAEEL 07A
> 365	95	14	ERLER	99	RVUE	Electroweak
> 87	95	15	CONRAD	98	RVUE	$\nu_\mu N$ scattering
> 620	95	16	ABE	97S	CDF	$p\bar{p}; Z'_\eta \rightarrow e^+e^-, \mu^+\mu^-$
> 100	95	17	VILAIN	94B	CHM2	$\nu_\mu e \rightarrow \nu_\mu e; \bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$
> 125	90	18	ABE	90F	VNS	$e^+e^-$
[> 820]		19	GONZALEZ-G.	90D	COSM	Nucleosynthesis; light $\nu_R$
[> 3300]		20	GRIFOLS	90	ASTR	SN 1987A; light $\nu_R$
[> 1040]		19	LOPEZ	90	COSM	Nucleosynthesis; light $\nu_R$

- 1 AAD 12CC search for resonances decaying to  $e^+e^-$ ,  $\mu^+\mu^-$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- 2 CHO 00 use various electroweak data to constrain  $Z'$  models assuming  $m_H = 100$  GeV. See Fig. 3 for limits in the mass-mixing plane.
- 3 AALTONEN 11I search for resonances decaying to  $\mu^+\mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 4 ABZOV 11A, AALTONEN 09T, AALTONEN 07H, and ABULENCIA 06L search for resonances decaying to  $e^+e^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 5 DEL-AGUILA 10 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0023 < \theta < 0.0027$ .
- 6 AALTONEN 09V search for resonances decaying to  $\mu^+\mu^-$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 7 ERLER 99 give 95% CL limit on the  $Z$ - $Z'$  mixing  $-0.0047 < \theta < 0.0021$ .
- 8 ABDALLAH 06C give 95% CL limit  $|\theta| < 0.0092$ . See their Fig. 14 for limit contours in the mass-mixing plane.
- 9 ABULENCIA 05A search for resonances decaying to electron or muon pairs in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.
- 10 ABBIENDI 04G give 95% CL limit on  $Z$ - $Z'$  mixing  $-0.00447 < \theta < 0.00331$ . See their Fig. 20 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 91$  to 207 GeV.
- 11 BARGER 03B limit is from the nucleosynthesis bound on the effective number of light neutrino  $\delta N_\nu < 1$ . The quark-hadron transition temperature  $T_C = 150$  MeV is assumed. The limit with  $T_C = 400$  MeV is  $> 3300$  GeV.
- 12 ABREU 00S give 95% CL limit on  $Z$ - $Z'$  mixing  $|\theta| < 0.0024$ . See their Fig. 6 for the limit contour in the mass-mixing plane.  $\sqrt{s} = 90$  to 189 GeV.
- 13 BARATE 00I search for deviations in cross section and asymmetries in  $e^+e^- \rightarrow$  fermions at  $\sqrt{s} = 90$  to 183 GeV. Assume  $\theta = 0$ . Bounds in the mass-mixing plane are shown in their Figure 18.
- 14 ERLER 99 give 90% CL limit on the  $Z$ - $Z'$  mixing  $-0.0062 < \theta < 0.0011$ .
- 15 CONRAD 98 limit is from measurements at CCFR, assuming no  $Z$ - $Z'$  mixing.
- 16 ABE 97S find  $\sigma(Z') \times B(e^+e^-, \mu^+\mu^-) < 40$  fb for  $m_{Z'} > 600$  GeV at  $\sqrt{s} = 1.8$  TeV.
- 17 VILAIN 94B assume  $m_t = 150$  GeV and  $\theta = 0$ . See Fig. 2 for limit contours in the mass-mixing plane.
- 18 ABE 90F use data for  $R$ ,  $R_{\ell\ell}$ , and  $A_{\ell\ell}$ . ABE 90F fix  $m_W = 80.49 \pm 0.43 \pm 0.24$  GeV and  $m_Z = 91.13 \pm 0.03$  GeV.
- 19 These authors claim that the nucleosynthesis bound on the effective number of light neutrinos ( $\delta N_\nu < 1$ ) constrains  $Z'$  masses if  $\nu_R$  is light ( $\lesssim 1$  MeV).
- 20 GRIFOLS 90 limit holds for  $m_{\nu_R} \lesssim 1$  MeV. See also RIZZO 91.

### Limits for other $Z'$

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 2400	95	1 KHACHATRY...16E	CMS	$Z' \rightarrow t\bar{t}$
		2 AAD	15AO ATLS	$Z' \rightarrow t\bar{t}$
		3 AAD	15AT ATLS	monotop
		4 AAD	15CD ATLS	$h \rightarrow Z'Z', Z'Z'; Z' \rightarrow \ell^+\ell^-$
		5 AAD	15O ATLS	$Z' \rightarrow e\mu, e\tau, \mu\tau$
		6 KHACHATRY...15F	CMS	monotop
		7 KHACHATRY...15O	CMS	$Z' \rightarrow hZ$
		8 AAD	14AT ATLS	$Z' \rightarrow Z\gamma$
		9 KHACHATRY...14A	RVUE	$Z' \rightarrow VV$
		10 MARTINEZ	14 RVUE	Electroweak
		11 AAD	13AI ATLS	$Z' \rightarrow e\mu, e\tau, \mu\tau$
		12 AAD	13AQ ATLS	$Z' \rightarrow t\bar{t}$
		13 AAD	13G ATLS	$Z' \rightarrow t\bar{t}$
		14 AALTONEN	13A CDF	$Z' \rightarrow t\bar{t}$
		15 CHATRCHYAN13AP	CMS	$Z' \rightarrow t\bar{t}$
		16 CHATRCHYAN13BM	CMS	$Z' \rightarrow t\bar{t}$
		17 AAD	12BV ATLS	$Z' \rightarrow t\bar{t}$
		18 AAD	12K ATLS	$Z' \rightarrow t\bar{t}$
		19 AALTONEN	12AR CDF	Chromophilic
		20 AALTONEN	12N CDF	$Z' \rightarrow \bar{t}u$
		21 ABZOV	12R D0	$Z' \rightarrow t\bar{t}$
		22 CHATRCHYAN12AI	CMS	$Z' \rightarrow t\bar{t}$
		23 CHATRCHYAN12AQ	CMS	$Z' \rightarrow t\bar{t}$
		24 CHATRCHYAN12BL	CMS	$Z' \rightarrow t\bar{t}$
		25 AAD	11H ATLS	$Z' \rightarrow e\mu$
		26 AAD	11Z ATLS	$Z' \rightarrow e\mu$
		27 AALTONEN	11AD CDF	$Z' \rightarrow t\bar{t}$
		28 AALTONEN	11AE CDF	$Z' \rightarrow t\bar{t}$
		29 CHATRCHYAN11O	CMS	$pp \rightarrow t\bar{t}$
		30 AALTONEN	08D CDF	$Z' \rightarrow t\bar{t}$

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

27	AALTONEN	08Y	CDF	$Z' \rightarrow t\bar{t}$	
27	ABAZOV	08AA	D0	$Z' \rightarrow t\bar{t}$	
28	ABULENCIA	06M	CDF	$Z' \rightarrow e\mu$	
29	ABAZOV	04A	D0	Repl. by ABAZOV 08AA	
30	BARGER	03B	COSM	Nucleosynthesis; light $\nu_R$	
31	CHO	00	RVUE	$E_6$ -motivated	
32	CHO	98	RVUE	$E_6$ -motivated	
33	ABE	97G	CDF	$Z' \rightarrow \bar{q}q$	

<sup>1</sup> KHACHATRYAN 16E search for a leptophobic top-color  $Z'$  decaying to  $t\bar{t}$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The quoted limit assumes that  $\Gamma_{Z'}/m_{Z'} = 0.012$ . Also  $m_{Z'} < 2.9$  TeV is excluded for wider topcolor  $Z'$  with  $\Gamma_{Z'}/m_{Z'} = 0.1$ .

<sup>2</sup> AAD 15A0 search for narrow resonance decaying to  $t\bar{t}$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. See Fig. 11 for limit on  $\sigma_B$ .

<sup>3</sup> AAD 15AT search for monoprot production plus large missing  $E_T$  events in  $pp$  collisions at  $\sqrt{s} = 8$  TeV and give constraints on a  $Z'$  model having  $Z'u\bar{t}$  coupling.  $Z'$  is assumed to decay invisibly. See their Fig. 6 for limits on  $\sigma \cdot B$ .

<sup>4</sup> AAD 15CD search for decays of Higgs bosons to 4  $\ell$  states via  $Z'$  bosons,  $h \rightarrow Z'Z' \rightarrow 4\ell$  or  $h \rightarrow Z'Z' \rightarrow 4\ell$ . See Fig. 5 for the limit on the signal strength of the  $h \rightarrow Z'Z' \rightarrow 4\ell$  process and Fig. 16 for the limit on  $h \rightarrow Z'Z' \rightarrow 4\ell$ .

<sup>5</sup> AAD 15o search for new particle with lepton flavor violating decay in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. See their Fig. 2 for limits on  $\sigma_B$ .

<sup>6</sup> KHACHATRYAN 15F search for monoprot production plus large missing  $E_T$  events in  $pp$  collisions at  $\sqrt{s} = 8$  TeV and give constraints on a  $Z'$  model having  $Z'u\bar{t}$  coupling.  $Z'$  is assumed to decay invisibly. See Fig. 3 for limits on  $\sigma_B$ .

<sup>7</sup> KHACHATRYAN 15O search for narrow  $Z'$  resonance decaying to  $Zh$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. See their Fig. 6 for limit on  $\sigma_B$ .

<sup>8</sup> AAD 14AT search for a narrow neutral vector boson decaying to  $Z\gamma$ . See their Fig. 3b for the exclusion limit in  $m_{Z'} - \sigma_B$  plane.

<sup>9</sup> KHACHATRYAN 14A search for new resonance in the  $WW$  ( $\ell\nu q\bar{q}$ ) and the  $ZZ$  ( $\ell\ell q\bar{q}$ ) channels using  $pp$  collisions at  $\sqrt{s}=8$  TeV. See their Fig.13 for the exclusion limit on the number of events in the mass-width plane.

<sup>10</sup> MARTINEZ 14 use various electroweak data to constrain the  $Z'$  boson in the 3-3-1 models.

<sup>11</sup> AAD 13A1 search for new particle with lepton flavor violating decay in  $pp$  collisions at  $\sqrt{s} = 7$  TeV. See their Fig. 2 for limits on  $\sigma \cdot B$ .

<sup>12</sup> AAD 13A0 search for a leptophobic top-color  $Z'$  decaying to  $t\bar{t}$ . The quoted limit assumes that  $\Gamma_{Z'}/m_{Z'} = 0.012$ .

<sup>13</sup> CHATRCHYAN 13BM search for top-color  $Z'$  decaying to  $t\bar{t}$  using  $pp$  collisions at  $\sqrt{s}=8$  TeV. The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ .

<sup>14</sup> CHATRCHYAN 13AP search for top-color leptophobic  $Z'$  decaying to  $t\bar{t}$  using  $pp$  collisions at  $\sqrt{s}=7$  TeV. The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ .

<sup>15</sup> AAD 12BV search for narrow resonance decaying to  $t\bar{t}$  using  $pp$  collisions at  $\sqrt{s}=7$  TeV. See their Fig. 7 for limit on  $\sigma \cdot B$ .

<sup>16</sup> AAD 12K search for narrow resonance decaying to  $t\bar{t}$  using  $pp$  collisions at  $\sqrt{s}=7$  TeV. See their Fig. 5 for limit on  $\sigma \cdot B$ .

<sup>17</sup> AALTONEN 12AR search for chromophilic  $Z'$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. See their Fig. 5 for limit on  $\sigma \cdot B$ .

<sup>18</sup> AALTONEN 12N search for  $p\bar{p} \rightarrow tZ', Z' \rightarrow \bar{t}u$  events in  $p\bar{p}$  collisions. See their Fig. 3 for the limit on  $\sigma \cdot B$ .

<sup>19</sup> ABAZOV 12R search for top-color  $Z'$  boson decaying exclusively to  $t\bar{t}$ . The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ .

<sup>20</sup> CHATRCHYAN 12AI search for  $pp \rightarrow t\bar{t}$  events and give constraints on a  $Z'$  model having  $Z'\bar{t}t$  coupling. See their Fig. 4 for the limit in mass-coupling plane.

<sup>21</sup> Search for resonance decaying to  $t\bar{t}$ . See their Fig. 6 for limit on  $\sigma \cdot B$ .

<sup>22</sup> AAD 11H search for new particle with lepton flavor violating decay in  $pp$  collisions at  $\sqrt{s} = 7$  TeV. See their Fig. 3 for exclusion plot on the production cross section.

<sup>23</sup> AAD 11Z search for new particle with lepton flavor violating decay in  $pp$  collisions at  $\sqrt{s} = 7$  TeV. See their Fig. 3 for limit on  $\sigma \cdot B$ .

<sup>24</sup> Search for narrow resonance decaying to  $t\bar{t}$ . See their Fig. 4 for limit on  $\sigma \cdot B$ .

<sup>25</sup> Search for narrow resonance decaying to  $t\bar{t}$ . See their Fig. 3 for limit on  $\sigma \cdot B$ .

<sup>26</sup> CHATRCHYAN 11o search for same-sign top production in  $pp$  collisions induced by a hypothetical FCNC  $Z'$  at  $\sqrt{s} = 7$  TeV. See their Fig. 3 for limit in mass-coupling plane.

<sup>27</sup> Search for narrow resonance decaying to  $t\bar{t}$ . See their Fig. 3 for limit on  $\sigma \cdot B$ .

<sup>28</sup> ABULENCIA 06M search for new particle with lepton flavor violating decay at  $\sqrt{s} = 1.96$  TeV. See their Fig.4 for an exclusion plot on a mass-coupling plane.

<sup>29</sup> Search for narrow resonance decaying to  $t\bar{t}$ . See their Fig. 2 for limit on  $\sigma \cdot B$ .

<sup>30</sup> BARGER 03b use the nucleosynthesis bound on the effective number of light neutrino  $\delta N_\nu$ . See their Figs. 4-5 for limits in general  $E_6$  motivated models.

<sup>31</sup> CHO 00 use various electroweak data to constrain  $Z'$  models assuming  $m_H=100$  GeV. See Fig. 2 for limits in general  $E_6$ -motivated models.

<sup>32</sup> CHO 98 study constraints on four-Fermi contact interactions obtained from low-energy electroweak experiments, assuming no  $Z$ - $Z'$  mixing.

<sup>33</sup> Search for  $Z'$  decaying to dijets at  $\sqrt{s}=1.8$  TeV. For  $Z'$  with electromagnetic strength coupling, no bound is obtained.

### Indirect Constraints on Kaluza-Klein Gauge Bosons

Bounds on a Kaluza-Klein excitation of the  $Z$  boson or photon in  $d=1$  extra dimension. These bounds can also be interpreted as a lower bound on  $1/R$ , the size of the extra dimension. Unless otherwise stated, bounds assume all fermions live on a single brane and all gauge fields occupy the 4+ $d$ -dimensional bulk. See also the section on "Extra Dimensions" in the "Searches" Listings in this Review.

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • •				We do not use the following data for averages, fits, limits, etc. • • •

> 4.7		1	MUECK	02	RVUE	Electroweak
> 3.3	95	2	CORNET	00	RVUE	$e\nu qq'$
>5000		3	DELGADO	00	RVUE	$\epsilon_K$
> 2.6	95	4	DELGADO	00	RVUE	Electroweak
> 3.3	95	5	RIZZO	00	RVUE	Electroweak
> 2.9	95	6	MARCIAÑO	99	RVUE	Electroweak
> 2.5	95	7	MASIP	99	RVUE	Electroweak
> 1.6	90	8	NATH	99	RVUE	Electroweak
> 3.4	95	9	STRUMIA	99	RVUE	Electroweak

- <sup>1</sup> MUECK 02 limit is  $2\sigma$  and is from global electroweak fit ignoring correlations among observables. Higgs is assumed to be confined on the brane and its mass is fixed. For scenarios of bulk Higgs, of brane-SU(2)<sub>L</sub>, bulk-U(1)<sub>Y</sub>, and of bulk-SU(2)<sub>L</sub>, brane-U(1)<sub>Y</sub>, the corresponding limits are  $> 4.6$  TeV,  $> 4.3$  TeV and  $> 3.0$  TeV, respectively.
- <sup>2</sup> Bound is derived from limits on  $e\nu qq'$  contact interaction, using data from HERA and the Tevatron.
- <sup>3</sup> Bound holds only if first two generations of quarks lives on separate branes. If quark mixing is not complex, then bound lowers to 400 TeV from  $\Delta m_K$ .
- <sup>4</sup> See Figs. 1 and 2 of DELGADO 00 for several model variations. Special boundary conditions can be found which permit KK states down to 950 GeV and that agree with the measurement of  $Q_W(\text{Cs})$ . Quoted bound assumes all Higgs bosons confined to brane; placing one Higgs doublet in the bulk lowers bound to 2.3 TeV.
- <sup>5</sup> Bound is derived from global electroweak analysis assuming the Higgs field is trapped on the matter brane. If the Higgs propagates in the bulk, the bound increases to 3.8 TeV.
- <sup>6</sup> Bound is derived from global electroweak analysis but considering only presence of the KK  $W$  bosons.
- <sup>7</sup> Global electroweak analysis used to obtain bound independent of position of Higgs on brane or in bulk.
- <sup>8</sup> Bounds from effect of KK states on  $G_F$ ,  $\alpha$ ,  $M_W$ , and  $M_Z$ . Hard cutoff at string scale determined using gauge coupling unification. Limits for  $d=2,3,4$  rise to 3.5, 5.7, and 7.8 TeV.
- <sup>9</sup> Bound obtained for Higgs confined to the matter brane with  $m_H=500$  GeV. For Higgs in the bulk, the bound increases to 3.5 TeV.

## LEPTOQUARKS

Updated September 2015 by S. Rolli (US Department of Energy) and M. Tanabashi (Nagoya U.)

Leptoquarks are hypothetical particles carrying both baryon number (B) and lepton number (L). The possible quantum numbers of leptoquark states can be restricted by assuming that their direct interactions with the ordinary SM fermions are dimensionless and invariant under the standard model (SM) gauge group. Table 1 shows the list of all possible quantum numbers with this assumption [1]. The columns of  $SU(3)_C$ ,  $SU(2)_W$ , and  $U(1)_Y$  in Table 1 indicate the QCD representation, the weak isospin representation, and the weak hypercharge, respectively. The spin of a leptoquark state is taken to be 1 (vector leptoquark) or 0 (scalar leptoquark).

**Table 1:** Possible leptoquarks and their quantum numbers.

Spin	$3B+L$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$	Allowed coupling
0	-2	$\bar{3}$	1	1/3	$\bar{q}_L^c \ell_L$ or $\bar{u}_R^c e_R$
0	-2	$\bar{3}$	1	4/3	$\bar{d}_R^c e_R$
0	-2	$\bar{3}$	3	1/3	$\bar{q}_L^c \ell_L$
1	-2	$\bar{3}$	2	5/6	$\bar{q}_L^c \gamma^\mu e_R$ or $\bar{d}_R^c \gamma^\mu \ell_L$
1	-2	$\bar{3}$	2	-1/6	$\bar{u}_R^c \gamma^\mu \ell_L$
0	0	3	2	7/6	$\bar{q}_L e_R$ or $\bar{u}_R \ell_L$
0	0	3	2	1/6	$\bar{d}_R \ell_L$
1	0	3	1	2/3	$\bar{q}_L \gamma^\mu \ell_L$ or $\bar{d}_R \gamma^\mu e_R$
1	0	3	1	5/3	$\bar{u}_R \gamma^\mu e_R$
1	0	3	3	2/3	$\bar{q}_L \gamma^\mu \ell_L$

If we do not require leptoquark states to couple directly with SM fermions, different assignments of quantum numbers become possible [2,3].

Leptoquark states are expected to exist in various extensions of SM. The Pati-Salam model [4] is an example predicting the existence of a leptoquark state. Vector leptoquark states also exist in grand unification theories based on  $SU(5)$  [5],  $SO(10)$  [6], which includes Pati-Salam color  $SU(4)$ , and larger gauge groups. Scalar quarks in supersymmetric models with R-parity violation may also have leptoquark-type Yukawa couplings. The bounds on the leptoquark states can therefore be applied to constrain R-parity-violating supersymmetric models. Scalar leptoquarks are expected to exist at TeV scale in extended technicolor models [7,8] where leptoquark states appear as the bound states of techni-fermions. Compositeness of quarks and leptons also provides examples of models which may have light leptoquark states [9].

Bounds on leptoquark states are obtained both directly and indirectly. Direct limits are from their production cross sections at colliders, while indirect limits are calculated from the bounds on the leptoquark-induced four-fermion interactions, which are obtained from low-energy experiments, or from collider experiments below threshold.

If a leptoquark couples to fermions belonging to more than a single generation in the mass eigenbasis of the SM fermions, it can induce four-fermion interactions causing flavor-changing neutral currents and lepton-family-number violations. The quantum number assignment of Table 1 allows several leptoquark states to couple to both left- and right-handed quarks simultaneously. Such leptoquark states are called non-chiral and may cause four-fermion interactions affecting the  $(\pi \rightarrow e\nu)/(\pi \rightarrow \mu\nu)$  ratio [10]. Non-chiral scalar leptoquarks also contribute to the muon anomalous magnetic moment [11,12]. Since indirect limits provide more stringent constraints on these types of leptoquarks, it is often assumed that a leptoquark state couples only to a single generation in a chiral interaction, for which indirect limits become much weaker. Additionally, this assumption gives strong constraints on concrete models of leptoquarks.

Leptoquark states which couple only to left- or right-handed quarks are called chiral leptoquarks. Leptoquark states which couple only to the first (second, third) generation are referred as the first- (second-, third-) generation leptoquarks. Refs. [13,14] give extensive lists of the bounds on the leptoquark-induced four-fermion interactions. For the isoscalar and vector leptoquarks  $S_0$  and  $V_0$ , for example, which couple with the first- (second-) generation left-handed quark, and the first-generation left-handed lepton, the bounds of Ref. 13 read  $\lambda^2 < 0.03 \times (M_{LQ}/300 \text{ GeV})^2$  for  $S_0$ , and  $\lambda^2 < 0.02 \times (M_{LQ}/300 \text{ GeV})^2$  for  $V_0$  ( $\lambda^2 < 5 \times (M_{LQ}/300 \text{ GeV})^2$  for  $S_0$ , and  $\lambda^2 < 3 \times (M_{LQ}/300 \text{ GeV})^2$  for  $V_0$ ) with  $\lambda$  being the leptoquark coupling strength. The  $e^+e^-$  experiments are sensitive to the indirect effects coming from  $t$ - and  $u$ -channel exchanges of leptoquarks in the  $e^+e^- \rightarrow q\bar{q}$  process.

The HERA experiments give bounds on the leptoquark-induced four-fermion interaction. For detailed bounds obtained in this way, see the Boson Particle Listings for “Indirect Limits for Leptoquarks” and its references.

Collider experiments provide direct limits on the leptoquark states through limits on the pair- and single-production cross sections. The leading-order cross sections of the parton processes

$$\begin{aligned} q + \bar{q} &\rightarrow LQ + \overline{LQ} \\ g + g &\rightarrow LQ + \overline{LQ} \\ e + q &\rightarrow LQ \end{aligned} \quad (1)$$

may be written as [15]

$$\begin{aligned} \hat{\sigma}_{LO}[q\bar{q} \rightarrow LQ + \overline{LQ}] &= \frac{2\alpha_s^2\pi}{27\hat{s}}\beta^3, \\ \hat{\sigma}_{LO}[gg \rightarrow LQ + \overline{LQ}] &= \frac{\alpha_s^2\pi}{96\hat{s}} \\ &\times \left[ \beta(41 - 31\beta^2) + (18\beta^2 - \beta^4 - 17) \log \frac{1+\beta}{1-\beta} \right], \\ \hat{\sigma}_{LO}[eq \rightarrow LQ] &= \frac{\pi\lambda^2}{4}\delta(\hat{s} - M_{LQ}^2) \end{aligned} \quad (2)$$

for a scalar leptoquark. Here  $\sqrt{\hat{s}}$  is the invariant energy of the parton subprocess, and  $\beta \equiv \sqrt{1 - 4M_{LQ}^2/\hat{s}}$ . The leptoquark Yukawa coupling is given by  $\lambda$ . Leptoquarks are also produced singly at hadron colliders through  $g + q \rightarrow LQ + \ell$  [16], which allows extending to higher masses the collider reach in the leptoquark search [17], depending on the leptoquark Yukawa coupling.

The LHC, Tevatron and LEP experiments search for pair production of the leptoquark states, which arises from the leptoquark gauge interaction. The searches are carried on in signatures including high  $P_T$  leptons,  $E_T$  jets and large missing transverse energy, due to the typical decay of the leptoquark. The gauge couplings of a scalar leptoquark are determined uniquely according to its quantum numbers in Table 1. Since all of the leptoquark states belong to color-triplet representation, the scalar leptoquark pair-production cross section at the Tevatron and LHC can be determined solely as a function of the leptoquark mass without making further assumptions. This is in contrast to the indirect or single-production limits, which give constraints in the leptoquark mass-coupling plane. For the first- and second-generation scalar leptoquark states with decaying branching fraction  $\beta = B(eq) = 1$  and  $\beta = B(\mu q) = 1$ , the CDF and DØ experiments obtain the lower bounds on the leptoquark mass  $> 236 \text{ GeV}$  (first generation, CDF) [18],  $> 299 \text{ GeV}$  (first generation, DØ) [19],  $> 226 \text{ GeV}$  (second generation, CDF) [20], and  $> 316 \text{ GeV}$  (second generation, DØ) [21] at 95% CL. Third generation leptoquark mass bounds come from the DØ experiment [22] which sets a limit at  $247 \text{ GeV}$  for a charge  $-1/3$  third generation scalar leptoquark, at 95% C.L.

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Recent results from the LHC proton-proton collider, running at a center of mass energy of 7 and 8 TeV, extend previous Tevatron mass limits for scalar leptoquarks to  $> 830$  GeV (first generation, CMS,  $\beta = 1$ ,  $\sqrt{s} = 7$  TeV) and  $> 640$  GeV (first generation, CMS,  $\beta = 0.5$ ,  $\sqrt{s} = 7$  TeV) [23];  $> 1050$  GeV (first generation, ATLAS,  $\beta = 1$ ,  $\sqrt{s} = 8$  TeV) and  $> 900$  GeV (first generation, ATLAS,  $\beta = 0.5$ ,  $\sqrt{s} = 8$  TeV) [24];  $> 1070$  GeV (second generation, CMS,  $\beta = 1$ ,  $\sqrt{s} = 7$  TeV) [25] and  $> 785$  GeV (second generation, CMS,  $\beta = 0.5$ ,  $\sqrt{s} = 7$  TeV) [25]; and  $> 1000$  GeV (second generation, ATLAS,  $\beta = 1$ ,  $\sqrt{s} = 8$  TeV) and  $> 850$  GeV (second generation, ATLAS,  $\beta = 0.5$ ,  $\sqrt{s} = 8$  TeV) [24]. All limits at 95% C.L.

As for third generation leptoquarks, CMS results are the following: 1) assuming that all leptoquarks decay to a top quark and a  $\tau$  lepton, the existence of pair produced, third-generation leptoquarks up to a mass of 685 GeV ( $\beta = 1$ ) is excluded at 95% confidence level [26]; 2) assuming that all leptoquarks decay to a bottom quark and a  $\tau$  lepton, the existence of pair produced, third-generation leptoquarks up to a mass of 740 GeV ( $\beta = 1$ ) is excluded at 95% confidence level [27]; 3) assuming that all leptoquarks decay to a bottom quark and a  $\tau$  neutrino, the existence of pair produced, third-generation leptoquarks up to a mass of 450 GeV ( $\beta = 0.5$ ) is excluded at 95% confidence level [28].

The ATLAS collaboration has a limit on third generation scalar leptoquark for the case of  $\beta = 1$  of 525 GeV [29] and 625 GeV for third-generation leptoquarks in the bottom  $\tau$  neutrino channel, and 640 GeV in the top  $\tau$  neutrino channel [24].

The magnetic-dipole-type and the electric-quadrupole-type interactions of a vector leptoquark are not determined even if we fix its gauge quantum numbers as listed in the Table [30]. The production of vector leptoquarks depends in general on additional assumptions that the leptoquark couplings and their pair-production cross sections are enhanced relative to the scalar leptoquark contributions. At the Tevatron for instance, since the acceptance for vector and scalar leptoquark detection is similar, limits on the vector leptoquark mass will be more stringent (see for example [36,19]). The leptoquark pair-production cross sections in  $e^+e^-$  collisions depend on the leptoquark  $SU(2) \times U(1)$  quantum numbers and Yukawa coupling with electron [31]. The OPAL experiment sets mass bounds on various leptoquark states from the pair-production cross sections [32]. For a second-generation weak-isosinglet weak-hypercharge  $-4/3$  scalar-leptoquark state, for example, the OPAL pair-production bound is  $M_{LQ} > 100$  GeV/ $c^2$  at 95% C.L. The LEP experiments also searched for the single production of the leptoquark states from the process  $e\gamma \rightarrow LQ + q$ .

The most stringent searches for the leptoquark single production are performed by the HERA experiments. Since the leptoquark single-production cross section depends on its Yukawa coupling, the leptoquark mass limits from HERA are usually displayed in the mass-coupling plane. For leptoquark Yukawa coupling  $\lambda = 0.1$ , the ZEUS bounds on the first-generation

leptoquarks range from 248 to 290 GeV, depending on the leptoquark species [33]. Recently the H1 Collaboration released a comprehensive summary of searches for first generation leptoquarks using the full data sample collected in  $ep$  collisions at HERA (446 pb $^{-1}$ ). No evidence of production of leptoquarks is observed in final states with a large transverse momentum electron or large missing transverse momentum. For a coupling strength  $\lambda = 0.3$ , first generation leptoquarks with masses up to 800 GeV are excluded at 95% C.L. [35]

The search for LQ will be continued with more LHC data. Early feasibility studies by the LHC experiments ATLAS [37] and CMS [38] indicate that clear signals can be established for masses up to about  $M(LQ)$  1.3 to 1.4 TeV for first- and second-generation scalar LQ, with a likely final reach 1.5 TeV, for collisions at 14 TeV in the center of mass.

## Reference

1. W. Buchmüller, R. Rückl, and D. Wyler, Phys. Lett. **B191**, 442 (1987).
2. K.S. Babu, C.F. Kolda, and J. March-Russell, Phys. Lett. **B408**, 261 (1997).
3. J.L. Hewett and T.G. Rizzo, Phys. Rev. **D58**, 055005 (1998).
4. J.C. Pati and A. Salam, Phys. Rev. **D10**, 275 (1974).
5. H. Georgi and S.L. Glashow, Phys. Rev. Lett. **32**, 438 (1974).
6. H. Georgi, AIP Conf. Proc. **23**, 575 (1975); H. Fritzsch and P. Minkowski, Ann. Phys. **93**, 193 (1975).
7. For a review, see, E. Farhi and L. Susskind, Phys. Reports **74**, 277 (1981).
8. K. Lane and M. Ramana, Phys. Rev. **D44**, 2678 (1991).
9. See, for example, B. Schrempp and F. Schrempp, Phys. Lett. **153B**, 101 (1985).
10. O. Shanker, Nucl. Phys. **B204**, 375, (1982).
11. U. Mahanta, Eur. Phys. J. **C21**, 171 (2001) [Phys. Lett. **B515**, 111 (2001)].
12. K. Cheung, Phys. Rev. **D64**, 033001 (2001).
13. S. Davidson, D.C. Bailey, and B.A. Campbell, Z. Phys. **C61**, 613 (1994).
14. M. Leurer, Phys. Rev. **D49**, 333 (1994); Phys. Rev. **D50**, 536 (1994).
15. T. Plehn *et al.*, Z. Phys. **C74**, 611 (1997); M. Kramer *et al.*, Phys. Rev. Lett. **79**, 341 (1997); and references therein.
16. J.L. Hewett and S. Pakvasa, Phys. Rev. **D37**, 3165 (1988); O.J.P. Eboli and A.V. Olinto, Phys. Rev. **D38**, 3461 (1988); A. Dobado, M.J. Herrero, and C. Muñoz, Phys. Lett. **207B**, 97 (1988); V.D. Barger *et al.*, Phys. Lett. **B220**, 464 (1989); M. De Montigny and L. Marleau, Phys. Rev. **D40**, 2869 (1989) [Erratum-*ibid.* **D56**, 3156 (1997)].
17. A. Belyaev *et al.*, JHEP **0509**, 005 (2005).
18. D. Acosta *et al.* [CDF Collab.], Phys. Rev. **D72**, 051107 (2005).
19. V.M. Abazov *et al.* [DØCollab.], Phys. Lett. **B681**, 224 (2009).



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20. A. Abulencia *et al.* [CDF Collab.], Phys. Rev. **D73**, 051102 (2006).
21. V.M. Abazov *et al.* [DØCollab.], Phys. Lett. **B671**, 224 (2009).
22. V. Abazov *et al.* [DØCollab.], Phys. Lett. **B693**, 95 (2010).
23. S. Chatrchyan *et al.* [CMS Collab.], Phys. Rev. **D86**, 052013 (2012).
24. G. Aad *et al.* [ATLAS Collab.], arXiv:1508.04735v1.
25. S. Chatrchyan *et al.* [CMS Collab.], CMS PAS EXO-12-042 (2013).
26. V. Khachatryan *et al.* [CMS Collab.], JHEP **07**, 042 (2015).
27. V. Khachatryan *et al.* [CMS Collab.], Phys. Lett. **B739**, 229 (2014).
28. S. Chatrchyan *et al.* [CMS Collab.], JHEP **012**, 055 (2012).
29. G. Aad *et al.* [ATLAS Collab.], JHEP **06**, 033 (2013).
30. J. Blümlein, E. Boos, and A. Kryukov, Z. Phys. **C76**, 137 (1997).
31. J. Blümlein and R. Ruckl, Phys. Lett. **B304**, 337 (1993).
32. G. Abbiendi *et al.* [OPAL Collab.], Eur. Phys. J. **C31**, 281 (2003).
33. S. Chekanov *et al.* [ZEUS Collab.], Phys. Rev. **D68**, 052004 (2003).
34. A. Aktas *et al.* [H1 Collab.], Phys. Lett. **B629**, 9 (2005).
35. F.D. Aaron *et al.* [H1 Collab.], Phys. Lett. **B704**, 388 (2011).
36. T. Aalton *et al.* [CDF Collab.], Phys. Rev. **D77**, 091105 (2008).
37. V.A. Mitsou *et al.*, Czech. J. Phys. **55**, B659 (2005).
38. S. Abdulin and F. Charles, Phys. Lett. **B464**, 223 (1999).

### MASS LIMITS for Leptoquarks from Pair Production

These limits rely only on the color or electroweak charge of the leptoquark.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>1050	95	1 AAD 16G ATLS	First generation	
>1000	95	2 AAD 16G ATLS	Second generation	
> 625	95	3 AAD 16G ATLS	Third generation	
none 200–640	95	4 AAD 16G ATLS	Third generation	
> 685	95	5 KHACHATRYAN 15AJ CMS	Third generation	
> 740	95	6 KHACHATRYAN 14T CMS	Third generation	
> 534	95	7 AAD 13AE ATLS	Third generation	
> 830	95	8 CHATRCHYAN 12AG CMS	First generation	
> 840	95	9 CHATRCHYAN 12AG CMS	Second generation	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 525	95	10 CHATRCHYAN 13M CMS	Third generation	
> 660	95	11 AAD 12H ATLS	First generation	
> 685	95	12 AAD 12O ATLS	Second generation	
> 450	95	13 CHATRCHYAN 12BO CMS	Third generation	
> 376	95	14 AAD 11D ATLS	Superseded by AAD 12H	
> 422	95	15 AAD 11D ATLS	Superseded by AAD 12O	
> 326	95	16 ABAZOV 11V D0	First generation	
> 339	95	17 CHATRCHYAN 11N CMS	Superseded by CHATRCHYAN 12AG	
> 384	95	18 KHACHATRYAN 11D CMS	Superseded by CHATRCHYAN 12AG	
> 394	95	19 KHACHATRYAN 11E CMS	Superseded by CHATRCHYAN 12AG	
> 247	95	20 ABAZOV 10L D0	Third generation	
> 316	95	21 ABAZOV 09 D0	Second generation	
> 299	95	22 ABAZOV 09AF D0	Superseded by ABAZOV 11V	
> 153	95	23 AALTONEN 08P CDF	Third generation	
> 205	95	24 AALTONEN 08Z CDF	Third generation	
> 210	95	25 ABAZOV 08AD D0	All generations	
> 229	95	26 ABAZOV 08AN D0	Third generation	
> 251	95	27 ABAZOV 06A D0	Superseded by ABAZOV 10L	
> 136	95	28 ABAZOV 06L D0	Superseded by ABAZOV 08AD	
> 226	95	29 ABULENCIA 06T CDF	Second generation	
> 256	95	30 ABAZOV 05H D0	First generation	
> 117	95	25 ACOSTA 05I CDF	First generation	
> 236	95	31 ACOSTA 05P CDF	First generation	
> 99	95	32 ABBIENDI 03R OPAL	First generation	

> 100	95	32 ABBIENDI 03R OPAL	Second generation
> 98	95	32 ABBIENDI 03R OPAL	Third generation
> 98	95	33 ABAZOV 02 D0	All generations
> 225	95	34 ABAZOV 01D D0	First generation
> 85.8	95	35 ABBIENDI 00M OPAL	Superseded by ABBIENDI 03R
> 85.5	95	35 ABBIENDI 00M OPAL	Superseded by ABBIENDI 03R
> 82.7	95	35 ABBIENDI 00M OPAL	Superseded by ABBIENDI 03R
> 200	95	36 ABBOTT 00C D0	Second generation
> 123	95	37 AFFOLDER 00K CDF	Second generation
> 148	95	38 AFFOLDER 00K CDF	Third generation
> 160	95	39 ABBOTT 99J D0	Second generation
> 225	95	40 ABBOTT 98E D0	First generation
> 94	95	41 ABBOTT 98J D0	Third generation
> 202	95	42 ABE 98S CDF	Second generation
> 242	95	43 GROSS-PILCH. 98	First generation
> 99	95	44 ABE 97F CDF	Third generation
> 213	95	45 ABE 97X CDF	First generation
> 45.5	95	46,47 ABREU 93J DLPH	First + second generation
> 44.4	95	48 ADRIANI 93M L3	First generation
> 44.5	95	48 ADRIANI 93M L3	Second generation
> 45	95	48 DECAMP 92 ALEP	Third generation
none 8.9–22.6	95	49 KIM 90 AMY	First generation
none 10.2–23.2	95	49 KIM 90 AMY	Second generation
none 5–20.8	95	50 BARTEL 87B JADE	
none 7–20.5	95	51 BEHREND 86B CELL	

1 AAD 16G search for scalar leptoquarks using  $eejj$  events in collisions at  $\sqrt{s} = 8$  TeV. The limit above assumes  $B(eq) = 1$ .

2 AAD 16G search for scalar leptoquarks using  $\mu\mu jj$  events in collisions at  $\sqrt{s} = 8$  TeV. The limit above assumes  $B(\mu q) = 1$ .

3 AAD 16G search for scalar leptoquarks decaying to  $b\nu$ . The limit above assumes  $B(b\nu) = 1$ .

4 AAD 16G search for scalar leptoquarks decaying to  $t\nu$ . The limit above assumes  $B(t\nu) = 1$ .

5 KHACHATRYAN 15AJ search for scalar leptoquarks using  $\tau\tau tt$  events in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The limit above assumes  $B(\tau\tau) = 1$ .

6 KHACHATRYAN 14T search for scalar leptoquarks decaying to  $\tau b$  using  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The limit above assumes  $B(\tau b) = 1$ . See their Fig. 5 for exclusion limit as function of  $B(\tau b)$ .

7 AAD 13AE search for scalar leptoquarks using  $\tau\tau bb$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(\tau b) = 1$ .

8 CHATRCHYAN 12AG search for scalar leptoquarks using  $eejj$  and  $e\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$ , the limit becomes 640 GeV.

9 CHATRCHYAN 12AG search for scalar leptoquarks using  $\mu\mu jj$  and  $\mu\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$ , the limit becomes 650 GeV.

10 CHATRCHYAN 13M search for scalar and vector leptoquarks decaying to  $\tau b$  in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above is for scalar leptoquarks with  $B(\tau b) = 1$ .

11 AAD 12H search for scalar leptoquarks using  $eejj$  and  $e\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$ , the limit becomes 607 GeV.

12 AAD 12O search for scalar leptoquarks using  $\mu\mu jj$  and  $\mu\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$ , the limit becomes 594 GeV.

13 CHATRCHYAN 12BO search for scalar leptoquarks decaying to  $\nu b$  in  $pp$  collisions at  $\sqrt{s} = 7$  TeV. The limit above assumes  $B(\nu b) = 1$ .

14 AAD 11D search for scalar leptoquarks using  $eejj$  and  $e\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$ , the limit becomes 319 GeV.

15 AAD 11D search for scalar leptoquarks using  $\mu\mu jj$  and  $\mu\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$ , the limit becomes 362 GeV.

16 ABAZOV 11V search for scalar leptoquarks using  $e\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(eq) = 0.5$ .

17 CHATRCHYAN 11N search for scalar leptoquarks using  $e\nu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(eq) = 0.5$ .

18 KHACHATRYAN 11D search for scalar leptoquarks using  $eejj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(eq) = 1$ .

19 KHACHATRYAN 11E search for scalar leptoquarks using  $\mu\mu jj$  events in  $pp$  collisions at  $E_{cm} = 7$  TeV. The limit above assumes  $B(\mu q) = 1$ .

20 ABAZOV 10L search for pair productions of scalar leptoquark state decaying to  $\nu b$  in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(\nu b) = 1$ .

21 ABAZOV 09 search for scalar leptoquarks using  $\mu\mu jj$  and  $\mu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$ , the limit becomes 270 GeV.

22 ABAZOV 09AF search for scalar leptoquarks using  $eejj$  and  $e\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$  the bound becomes 284 GeV.

23 AALTONEN 08P search for vector leptoquarks using  $\tau^+ b \bar{b}$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. Assuming Yang-Mills (minimal) couplings, the mass limit is  $>317$  GeV (251 GeV) at 95% CL for  $B(\tau b) = 1$ .

24 Search for pair production of scalar leptoquark state decaying to  $\tau b$  in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(\tau b) = 1$ .

25 Search for scalar leptoquarks using  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(\nu q) = 1$ .

26 ABAZOV 07J search for pair productions of scalar leptoquark state decaying to  $\nu b$  in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(\nu b) = 1$ .

27 ABAZOV 06A search for scalar leptoquarks using  $\mu\mu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV and 1.96 TeV. The limit above assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$ , the limit becomes 204 GeV.

28 ABAZOV 06L search for scalar leptoquarks using  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV and at 1.96 TeV. The limit above assumes  $B(\nu q) = 1$ .

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- <sup>29</sup> ABULENCIA 06T search for scalar leptoquarks using  $\mu\mu jj$ ,  $\nu\nu jj$ , and  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The quoted limit assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$  or  $0.1$ , the bound becomes 208 GeV or 143 GeV, respectively. See their Fig. 4 for the exclusion limit as a function of  $B(\mu q)$ .
- <sup>30</sup> ABAZOV 05H search for scalar leptoquarks using  $eejj$  and  $e\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV and 1.96 TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$  the bound becomes 234 GeV.
- <sup>31</sup> ACOSTA 05P search for scalar leptoquarks using  $eejj$ ,  $e\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.96$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$  and  $0.1$ , the bound becomes 205 GeV and 145 GeV, respectively.
- <sup>32</sup> ABBIENDI 03R search for scalar/vector leptoquarks in  $e^+e^-$  collisions at  $\sqrt{s} = 189$ –209 GeV. The quoted limits are for charge  $-4/3$  isospin 0 scalar-leptoquark with  $B(\ell q) = 1$ . See their table 12 for other cases.
- <sup>33</sup> ABAZOV 02 search for scalar leptoquarks using  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The bound holds for all leptoquark generations. Vector leptoquarks are likewise constrained to lie above 200 GeV.
- <sup>34</sup> ABAZOV 01D search for scalar leptoquarks using  $e\nu jj$ ,  $eejj$ , and  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$  and  $0$ , the bound becomes 204 and 79 GeV, respectively. Bounds for vector leptoquarks are also given. Supersedes ABBOTT 98E.
- <sup>35</sup> ABBIENDI 00M search for scalar/vector leptoquarks in  $e^+e^-$  collisions at  $\sqrt{s} = 183$  GeV. The quoted limits are for charge  $-4/3$  isospin 0 scalar-leptoquarks with  $B(\ell q) = 1$ . See their Table 8 and Figs. 6–9 for other cases.
- <sup>36</sup> ABBOTT 00C search for scalar leptoquarks using  $\mu\mu jj$ ,  $\nu\nu jj$ , and  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The limit above assumes  $B(\mu q) = 1$ . For  $B(\mu q) = 0.5$  and  $0$ , the bound becomes 180 and 79 GeV respectively. Bounds for vector leptoquarks are also given.
- <sup>37</sup> AFFOLDER 00K search for scalar leptoquark using  $\nu\nu cc$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The quoted limit assumes  $B(\nu c) = 1$ . Bounds for vector leptoquarks are also given.
- <sup>38</sup> AFFOLDER 00K search for scalar leptoquark using  $\nu\nu bb$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The quoted limit assumes  $B(\nu b) = 1$ . Bounds for vector leptoquarks are also given.
- <sup>39</sup> ABBOTT 99J search for leptoquarks using  $\mu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The quoted limit is for a scalar leptoquark with  $B(\mu q) = B(\nu q) = 0.5$ . Limits on vector leptoquarks range from 240 to 290 GeV.
- <sup>40</sup> ABBOTT 98E search for scalar leptoquarks using  $e\nu jj$ ,  $eejj$ , and  $\nu\nu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The limit above assumes  $B(eq) = 1$ . For  $B(eq) = 0.5$  and  $0$ , the bound becomes 204 and 79 GeV, respectively.
- <sup>41</sup> ABBOTT 98J search for charge  $-1/3$  third generation scalar and vector leptoquarks in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The quoted limit is for scalar leptoquark with  $B(\nu b) = 1$ .
- <sup>42</sup> ABE 98S search for scalar leptoquarks using  $\mu\mu jj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The limit is for  $B(\mu q) = 1$ . For  $B(\mu q) = B(\nu q) = 0.5$ , the limit is  $> 160$  GeV.
- <sup>43</sup> GROSS-PILCHER 98 is the combined limit of the CDF and DØ Collaborations as determined by a joint CDF/DØ working group and reported in this FNAL Technical Memo. Original data published in ABE 97X and ABBOTT 98E.
- <sup>44</sup> ABE 97F search for third generation scalar and vector leptoquarks in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The quoted limit is for scalar leptoquark with  $B(\tau b) = 1$ .
- <sup>45</sup> ABE 97X search for scalar leptoquarks using  $eejj$  events in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. The limit is for  $B(eq) = 1$ .
- <sup>46</sup> Limit is for charge  $-1/3$  isospin-0 leptoquark with  $B(\ell q) = 2/3$ .
- <sup>47</sup> First and second generation leptoquarks are assumed to be degenerate. The limit is slightly lower for each generation.
- <sup>48</sup> Limits are for charge  $-1/3$ , isospin-0 scalar leptoquarks decaying to  $\ell^- q$  or  $\nu q$  with any branching ratio. See paper for limits for other charge-isospin assignments of leptoquarks.
- <sup>49</sup> KIM 90 assume pair production of charge  $2/3$  scalar-leptoquark via photon exchange. The decay of the first (second) generation leptoquark is assumed to be any mixture of  $d e^+$  and  $u \bar{\nu}$  ( $s \mu^+$  and  $c \tau$ ). See paper for limits for specific branching ratios.
- <sup>50</sup> BARTEL 87B limit is valid when a pair of charge  $2/3$  spinless leptoquarks  $X$  is produced with point coupling, and when they decay under the constraint  $B(X \rightarrow c \bar{\nu}_\mu) + B(X \rightarrow s \mu^+) = 1$ .
- <sup>51</sup> BEHREND 86B assumed that a charge  $2/3$  spinless leptoquark,  $\chi$ , decays either into  $s \mu^+$  or  $c \tau$ .  $B(X \rightarrow s \mu^+) + B(X \rightarrow c \tau) = 1$ .

### MASS LIMITS for Leptoquarks from Single Production

These limits depend on the  $q$ - $\ell$ -leptoquark coupling  $g_{LQ}$ . It is often assumed that  $g_{LQ}^2/4\pi = 1/137$ . Limits shown are for a scalar, weak isoscalar, charge  $-1/3$  leptoquark.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>304	95	1 ABRAWOWICZ12A	ZEUS	First generation
> 73	95	2 ABREU 93J	DLPH	Second generation
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>300	95	3 AARON 11A	H1	Lepton-flavor violation
		4 AARON 11B	H1	First generation
		5 ABAZOV 07E	D0	Second generation
>295	95	6 AKTAS 05B	H1	First generation
		7 CHEKANOV 05A	ZEUS	Lepton-flavor violation
>298	95	8 CHEKANOV 03B	ZEUS	First generation
>197	95	9 ABBIENDI 02B	OPAL	First generation
		10 CHEKANOV 02	ZEUS	Repl. by CHEKANOV 05A
>290	95	11 ADLOFF 01C	H1	First generation
>204	95	12 BREITWEG 01	ZEUS	First generation
		13 BREITWEG 00E	ZEUS	First generation
>161	95	14 ABREU 99G	DLPH	First generation
>200	95	15 ADLOFF 99	H1	First generation
		16 DERRICK 97	ZEUS	Lepton-flavor violation
>168	95	17 DERRICK 93	ZEUS	First generation

- <sup>1</sup> ABRAWOWICZ 12A limit is for a scalar, weak isoscalar, charge  $-1/3$  leptoquark coupled with  $e\bar{\nu}$ . See their Figs. 12–17 and Table 4 for states with different quantum numbers.
- <sup>2</sup> Limit from single production in  $Z$  decay. The limit is for a leptoquark coupling of electromagnetic strength and assumes  $B(\ell q) = 2/3$ . The limit is 77 GeV if first and second leptoquarks are degenerate.
- <sup>3</sup> AARON 11A search for various leptoquarks with lepton-flavor violating couplings. See their Figs. 2–3 and Tables 1–4 for detailed limits.
- <sup>4</sup> The quoted limit is for a scalar, weak isoscalar, charge  $-1/3$  leptoquark coupled with  $e\bar{\nu}$ . See their Figs. 3–5 for limits on states with different quantum numbers.
- <sup>5</sup> ABAZOV 07E search for leptoquark single production through  $q\bar{q}$  fusion process in  $p\bar{p}$  collisions. See their Fig. 4 for exclusion plot in mass-coupling plane.
- <sup>6</sup> AKTAS 05B limit is for a scalar, weak isoscalar, charge  $-1/3$  leptoquark coupled with  $e\bar{\nu}$ . See their Fig. 3 for limits on states with different quantum numbers.
- <sup>7</sup> CHEKANOV 05 search for various leptoquarks with lepton-flavor violating couplings. See their Figs. 6–10 and Tables 1–8 for detailed limits.
- <sup>8</sup> CHEKANOV 03B limit is for a scalar, weak isoscalar, charge  $-1/3$  leptoquark coupled with  $e\bar{\nu}$ . See their Figs. 11–12 and Table 5 for limits on states with different quantum numbers.
- <sup>9</sup> For limits on states with different quantum numbers and the limits in the mass-coupling plane, see their Fig. 4 and Fig. 5.
- <sup>10</sup> CHEKANOV 02 search for various leptoquarks with lepton-flavor violating couplings. See their Figs. 6–7 and Tables 5–6 for detailed limits.
- <sup>11</sup> For limits on states with different quantum numbers and the limits in the mass-coupling plane, see their Fig. 3.
- <sup>12</sup> See their Fig. 14 for limits in the mass-coupling plane.
- <sup>13</sup> BREITWEG 00E search for  $F=0$  leptoquarks in  $e^+p$  collisions. For limits in mass-coupling plane, see their Fig. 11.
- <sup>14</sup> ABREU 99G limit obtained from process  $e\gamma \rightarrow LQ + q$ . For limits on vector and scalar states with different quantum numbers and the limits in the coupling-mass plane, see their Fig. 4 and Table 2.
- <sup>15</sup> For limits on states with different quantum numbers and the limits in the mass-coupling plane, see their Fig. 13 and Fig. 14. ADLOFF 99 also search for leptoquarks with lepton-flavor violating couplings. ADLOFF 99 supersedes AID 96B.
- <sup>16</sup> DERRICK 97 search for various leptoquarks with lepton-flavor violating couplings. See their Figs. 5–8 and Table 1 for detailed limits.
- <sup>17</sup> DERRICK 93 search for single leptoquark production in  $e p$  collisions with the decay  $e q$  and  $\nu q$ . The limit is for leptoquark coupling of electromagnetic strength and assumes  $B(e q) = B(\nu q) = 1/2$ . The limit for  $B(e q) = 1$  is 176 GeV. For limits on states with different quantum numbers, see their Table 3.

### Indirect Limits for Leptoquarks

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 14	95	1 BESSAA	15	RVUE $q\bar{q} \rightarrow e^+e^-$
		2 SAHOO	15A	RVUE $B_{s,d} \rightarrow \mu^+\mu^-$
		3 SAKAKI	13	RVUE $B \rightarrow D^{(*)}\tau\bar{\nu}$ , $B \rightarrow X_s\nu\bar{\nu}$
> 2.5	95	4 KOSNIK	12	RVUE $b \rightarrow s\ell^+\ell^-$
		5 AARON	11C	H1 First generation
		6 DORSNER	11	RVUE scalar, weak singlet, charge 4/3
> 0.49	95	7 AKTAS	07A	H1 Lepton-flavor violation
		8 SCHAEEL	07A	ALEP $e^+e^- \rightarrow q\bar{q}$
		9 SMIRNOV	07	RVUE $K \rightarrow e\mu$ , $B \rightarrow e\tau$
> 1.7	96	10 CHEKANOV	05A	ZEUS Lepton-flavor violation
		11 ADLOFF	03	H1 First generation
		12 CHANG	03	BELL Pati-Salam type
> 46	90	13 CHEKANOV	02	ZEUS Repl. by CHEKANOV 05A
		14 CHEUNG	01B	RVUE First generation
		15 ACCIARRI	00P	L3 $e^+e^- \rightarrow q\bar{q}$
> 0.39	95	16 ADLOFF	00	H1 First generation
		17 BARATE	00I	ALEP Repl. by SCHAEEL 07A
		18 BARGER	00	RVUE Cs
> 0.74	95	19 GABRIELLI	00	RVUE Lepton flavor violation
		20 ZARNECKI	00	RVUE $S_1$ leptoquark
		21 ABBIENDI	99	OPAL
> 19.3	95	22 ABE	98V	CDF $B_s \rightarrow e^\pm\mu^\mp$ , Pati-Salam type
		23 ACCIARRI	98J	L3 $e^+e^- \rightarrow q\bar{q}$
		24 ACKERSTAFF	98V	OPAL $e^+e^- \rightarrow q\bar{q}$ , $e^+e^- \rightarrow b\bar{b}$
> 0.76	95	25 DEANDREA	97	RVUE $\tilde{R}_2$ leptoquark
		26 DERRICK	97	ZEUS Lepton-flavor violation
		27 GROSSMAN	97	RVUE $B \rightarrow \tau^+\tau^- (X)$
>1200	95	28 JADACH	97	RVUE $e^+e^- \rightarrow q\bar{q}$
		29 KUZNETSOV	95B	RVUE Pati-Salam type
		30 MIZUKOSHI	95	RVUE Third generation scalar leptoquark
> 0.3	95	31 BHATTACH...	94	RVUE Spin-0 leptoquark coupled to $\bar{\nu}_R t_L$
		32 DAVIDSON	94	RVUE
		33 KUZNETSOV	94	RVUE Pati-Salam type
> 18	95	34 LEURER	94	RVUE First generation spin-1 leptoquark
		35 MAHANTA	94B	RVUE First generation spin-0 leptoquark
		36 SHANKER	82	RVUE $P$ and $T$ violation
> 1	95	37 SHANKER	82	RVUE Nonchiral spin-0 leptoquark
		38 SHANKER	82	RVUE Nonchiral spin-1 leptoquark
		39 SHANKER	82	RVUE

- <sup>1</sup> BESSAA 15 obtain limit on leptoquark induced four-fermion interactions from the ATLAS and CMS limit on the  $q\bar{q}e\bar{e}$  contact interactions.
- <sup>2</sup> SAHOO 15A obtain limit on leptoquark induced four-fermion interactions from  $B_{s,d} \rightarrow \mu^+\mu^-$  for  $\lambda \simeq O(1)$ .
- <sup>3</sup> SAKAKI 13 explain the  $B \rightarrow D^{(*)}\tau\bar{\nu}$  anomaly using Wilson coefficients of leptoquark-induced four-fermion operators.

See key on page 601

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

- 4 KOSNIK 12 obtains limits on leptoquark induced four-fermion interactions from  $b \rightarrow s \ell^+ \ell^-$  decays.
- 5 AARON 11c limit is for weak isotriplet spin-0 leptoquark at strong coupling  $\lambda = \sqrt{4\pi}$ . For the limits of leptoquarks with different quantum numbers, see their Table 3. Limits are derived from bounds of  $e q$  contact interactions.
- 6 DORSNER 11 give bounds on scalar, weak singlet, charge 4/3 leptoquark from  $K, B, \tau$  decays, meson mixings,  $LFV$ ,  $g-2$  and  $Z \rightarrow b\bar{b}$ .
- 7 AKTAS 07a search for lepton-flavor violation in  $ep$  collision. See their Tables 4–7 for limits on lepton-flavor violating four-fermion interactions induced by various leptoquarks.
- 8 SCHAE 07a limit is for the weak-isoscalar spin-0 left-handed leptoquark with the coupling of electromagnetic strength. For the limits of leptoquarks with different quantum numbers, see their Table 35.
- 9 SMIRNOV 07 obtains mass limits for the vector and scalar chiral leptoquark states from  $K \rightarrow e\mu, B \rightarrow e\tau$  decays.
- 10 CHEKANOV 05 search for various leptoquarks with lepton-flavor violating couplings. See their Figs.6–10 and Tables 1–8 for detailed limits.
- 11 ADLOFF 03 limit is for the weak isotriplet spin-0 leptoquark at strong coupling  $\lambda = \sqrt{4\pi}$ . For the limits of leptoquarks with different quantum numbers, see their Table 3. Limits are derived from bounds on  $e^\pm q$  contact interactions.
- 12 The bound is derived from  $B(B^0 \rightarrow e^\pm \mu^\mp) < 1.7 \times 10^{-7}$ .
- 13 CHEKANOV 02 search for lepton-flavor violation in  $ep$  collisions. See their Tables 1–4 for limits on lepton-flavor violating and four-fermion interactions induced by various leptoquarks.
- 14 CHEUNG 01b quoted limit is for a scalar, weak isoscalar, charge  $-1/3$  leptoquark with a coupling of electromagnetic strength. The limit is derived from bounds on contact interactions in a global electroweak analysis. For the limits of leptoquarks with different quantum numbers, see Table 5.
- 15 ACCIARRI 00p limit is for the weak isoscalar spin-0 leptoquark with the coupling of electromagnetic strength. For the limits of leptoquarks with different quantum numbers, see their Table 4.
- 16 ADLOFF 00 limit is for the weak isotriplet spin-0 leptoquark at strong coupling,  $\lambda = \sqrt{4\pi}$ . For the limits of leptoquarks with different quantum numbers, see their Table 2.
- 17 ADLOFF 00 limits are from the  $Q^2$  spectrum measurement of  $e^+p \rightarrow e^+X$ .
- 18 BARATE 00i search for deviations in cross section and jet-charge asymmetry in  $e^+e^- \rightarrow q\bar{q}$  due to  $t$ -channel exchange of a leptoquark at  $\sqrt{s}=130$  to 183 GeV. Limits for other scalar and vector leptoquarks are also given in their Table 22.
- 19 BARGER 00 explain the deviation of atomic parity violation in cesium atoms from prediction is explained by scalar leptoquark exchange.
- 20 GABRIELLI 00 calculate various process with lepton flavor violation in leptoquark models.
- 21 ZARNECKI 00 limit is derived from data of HERA, LEP, and Tevatron and from various low-energy data including atomic parity violation. Leptoquark coupling with electromagnetic strength is assumed.
- 22 ABBENDI 99 limits are from  $e^+e^- \rightarrow q\bar{q}$  cross section at 130–136, 161–172, 183 GeV. See their Fig. 8 and Fig. 9 for limits in mass-coupling plane.
- 23 ABE 98v quoted limit is from  $B(B_s \rightarrow e^\pm \mu^\mp) < 8.2 \times 10^{-6}$ . ABE 98v also obtain a similar limit on  $M_{LQ} > 20.4$  TeV from  $B(B_d \rightarrow e^\pm \mu^\mp) < 4.5 \times 10^{-6}$ . Both bounds assume the non-canonical association of the  $b$  quark with electrons or muons under SU(4).
- 24 ACCIARRI 98j limit is from  $e^+e^- \rightarrow q\bar{q}$  cross section at  $\sqrt{s}=130$ –172 GeV which can be affected by the  $t$ - and  $u$ -channel exchanges of leptoquarks. See their Fig. 4 and Fig. 5 for limits in the mass-coupling plane.
- 25 ACKERSTAFF 98v limits are from  $e^+e^- \rightarrow q\bar{q}$  and  $e^+e^- \rightarrow b\bar{b}$  cross sections at  $\sqrt{s}=130$ –172 GeV, which can be affected by the  $t$ - and  $u$ -channel exchanges of leptoquarks. See their Fig. 21 and Fig. 22 for limits of leptoquarks in mass-coupling plane.
- 26 DEANDREA 97 limit is for  $R_2$  leptoquark obtained from atomic parity violation (APV). The coupling of leptoquark is assumed to be electromagnetic strength. See Table 2 for limits of the four-fermion interactions induced by various scalar leptoquark exchange. DEANDREA 97 combines APV limit and limits from Tevatron and HERA. See Fig. 1–4 for combined limits of leptoquark in mass-coupling plane.
- 27 DERRICK 97 search for lepton-flavor violation in  $ep$  collision. See their Tables 2–5 for limits on lepton-flavor violating four-fermion interactions induced by various leptoquarks.
- 28 GROSSMAN 97 estimate the upper bounds on the branching fraction  $B \rightarrow \tau^+ \tau^- (X)$  from the absence of the  $B$  decay with large missing energy. These bounds can be used to constrain leptoquark induced four-fermion interactions.
- 29 JADACH 97 limit is from  $e^+e^- \rightarrow q\bar{q}$  cross section at  $\sqrt{s}=172.3$  GeV which can be affected by the  $t$ - and  $u$ -channel exchanges of leptoquarks. See their Fig. 1 for limits on vector leptoquarks in mass-coupling plane.
- 30 KUZNETSOV 95b use  $\pi, K, B, \tau$  decays and  $\mu e$  conversion and give a list of bounds on the leptoquark mass and the fermion mixing matrix in the Pati-Salam model. The quoted limit is from  $K_L \rightarrow \mu e$  decay assuming zero mixing.
- 31 MIZUKOSHI 95 calculate the one-loop radiative correction to the Z-physics parameters in various scalar leptoquark models. See their Fig. 4 for the exclusion plot of third generation leptoquark models in mass-coupling plane.
- 32 BHATTACHARYYA 94 limit is from one-loop radiative correction to the leptonic decay width of the  $Z$ .  $m_H=250$  GeV,  $\alpha_s(m_Z)=0.12$ ,  $m_t=180$  GeV, and the electroweak strength of leptoquark coupling are assumed. For leptoquark coupled to  $\bar{e}_L t_R, \bar{\nu}_L \tau_R$ , and  $\bar{\tau}_L$ , see Fig. 2 in BHATTACHARYYA 94b erratum and Fig. 3.
- 33 DAVIDSON 94 gives an extensive list of the bounds on leptoquark-induced four-fermion interactions from  $\pi, K, D, B, \mu, \tau$  decays and meson mixings, etc. See Table 15 of DAVIDSON 94 for detail.
- 34 KUZNETSOV 94 gives mixing independent bound of the Pati-Salam leptoquark from the cosmological limit on  $\pi^0 \rightarrow \nu\nu$ .
- 35 LEURER 94, LEURER 94b limits are obtained from atomic parity violation and apply to any chiral leptoquark which couples to the first generation with electromagnetic strength. For a nonchiral leptoquark, universality in  $\pi_{e2}$  decay provides a much more stringent bound.
- 36 MAHANTA 94 gives bounds of  $P$ - and  $T$ -violating scalar-leptoquark couplings from atomic and molecular experiments.
- 37 From  $(\pi \rightarrow e\nu)/(\pi \rightarrow \mu\nu)$  ratio. SHANKER 82 assumes the leptoquark induced four-fermion coupling  $4g^2/M^2 (\bar{\nu}_L e_L u_R) (\bar{d}_L e_R)$  with  $g=0.004$  for spin-0 leptoquark and  $g^2/M^2 (\bar{\nu}_L \gamma_\mu u_L) (\bar{d}_R \gamma^\mu e_R)$  with  $g \approx 0.6$  for spin-1 leptoquark.

### MASS LIMITS for Diquarks

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;4700 (CL = 95%) OUR LIMIT</b>				
none 1200–4700	95	<sup>1</sup> KHACHATRY...15v	CMS	$E_6$ diquark

• • • We do not use the following data for averages, fits, limits, etc. • • •

>3750	95	<sup>2</sup> CHATRCHYAN 13A	CMS	$E_6$ diquark
none 1000–4280	95	<sup>3</sup> CHATRCHYAN 13As	CMS	Superseded by KHACHATRYAN 15v
>3520	95	<sup>4</sup> CHATRCHYAN 11v	CMS	Superseded by CHATRCHYAN 13A
none 970–1080,	95	<sup>5</sup> KHACHATRY...10	CMS	Superseded by CHATRCHYAN 13A
1450–1600	95	<sup>6</sup> AALTONEN	09ac CDF	$E_6$ diquark
none 290–630	95	<sup>7</sup> ABE	97g CDF	$E_6$ diquark
none 290–420	95	<sup>8</sup> ABREU	94o DLPH	SUSY $E_6$ diquark
none 15–31.7	95			

- <sup>1</sup> KHACHATRYAN 15v search for resonances decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>2</sup> CHATRCHYAN 13A search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- <sup>3</sup> CHATRCHYAN 13As search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>4</sup> CHATRCHYAN 11v search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- <sup>5</sup> KHACHATRYAN 10 search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- <sup>6</sup> AALTONEN 09ac search for new narrow resonance decaying to dijets.
- <sup>7</sup> ABE 97g search for new particle decaying to dijets.
- <sup>8</sup> ABREU 94o limit is from  $e^+e^- \rightarrow \bar{c}sc$ . Range extends up to 43 GeV if diquarks are degenerate in mass.

### MASS LIMITS for $g_A$ (axigluon) and Other Color-Octet Gauge Bosons

Axigluons are massive color-octet gauge bosons in chiral color models and have axial-vector coupling to quarks with the same coupling strength as gluons.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;3600 (CL = 95%) OUR LIMIT</b>				
none 1300–3600	95	<sup>1</sup> KHACHATRY...15v	CMS	$pp \rightarrow g_A X, g_A \rightarrow 2j$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>2800	95	<sup>2</sup> KHACHATRY...16e	CMS	$pp \rightarrow \bar{t} t K K^* X, g_{KK^*} \rightarrow \bar{t} t$
		<sup>3</sup> KHACHATRY...15av	CMS	$pp \rightarrow \theta^0 \theta^0 \rightarrow b\bar{b} Z g$
		<sup>4</sup> AALTONEN 13r	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow \sigma\sigma, \sigma \rightarrow 2j$
>3360	95	<sup>5</sup> CHATRCHYAN 13A	CMS	$pp \rightarrow g_A X, g_A \rightarrow 2j$
none 1000–3270	95	<sup>6</sup> CHATRCHYAN 13As	CMS	Superseded by KHACHATRYAN 15v
none 250–740	95	<sup>7</sup> CHATRCHYAN 13Au	CMS	$pp \rightarrow 2g_A X, g_A \rightarrow 2j$
> 775	95	<sup>8</sup> ABZOV	12r D0	$p\bar{p} \rightarrow g_A X, g_A \rightarrow t\bar{t}$
>2470	95	<sup>9</sup> CHATRCHYAN 11v	CMS	Superseded by CHATRCHYAN 13A
none 1470–1520	95	<sup>10</sup> AALTONEN 10L	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow t\bar{t}$
		<sup>11</sup> KHACHATRY...10	CMS	Superseded by CHATRCHYAN 13A
none 260–1250	95	<sup>12</sup> AALTONEN 09ac	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow 2j$
> 910	95	<sup>13</sup> CHOUDHURY 07	RVUE	$p\bar{p} \rightarrow t\bar{t} X$
> 365	95	<sup>14</sup> DONCHESKI 98	RVUE	$\Gamma(Z \rightarrow \text{hadron})$
none 200–980	95	<sup>15</sup> ABE 97g	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow 2j$
none 200–870	95	<sup>16</sup> ABE 95N	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow q\bar{q}$
none 240–640	95	<sup>17</sup> ABE 93g	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow 2j$
> 50	95	<sup>18</sup> CUYPERS 91	RVUE	$\sigma(e^+e^- \rightarrow \text{hadrons})$
none 120–210	95	<sup>19</sup> ABE 90H	CDF	$p\bar{p} \rightarrow g_A X, g_A \rightarrow 2j$
> 29		<sup>20</sup> ROBINETT 89	THEO	Partial-wave unitarity
none 150–310	95	<sup>21</sup> ALBAJAR 88b	UA1	$p\bar{p} \rightarrow g_A X, g_A \rightarrow 2j$
> 20		<sup>22</sup> BERGSTROM 88	RVUE	$p\bar{p} \rightarrow T X$ via $g_A g$
> 9		<sup>23</sup> CUYPERS 88	RVUE	$T$ decay
> 25		<sup>24</sup> DONCHESKI 88b	RVUE	$T$ decay

- <sup>1</sup> KHACHATRYAN 15v search for resonances decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>2</sup> KHACHATRYAN 16e search for KK gluon decaying to  $t\bar{t}$  in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>3</sup> KHACHATRYAN 15av search for pair productions of neutral color-octet weak-triplet scalar particles ( $\theta^0$ ), decaying to  $b\bar{b}, Z g$  or  $\gamma g$ , in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The  $\theta^0$  particle is often predicted in coloron ( $G'$ , color-octet gauge boson) models and appear in the  $pp$  collisions through  $G' \rightarrow \theta^0 \theta^0$  decays. Assuming  $B(\theta^0 \rightarrow b\bar{b}) = 0.5$ , they give limits  $m_{\theta^0} > 623$  GeV (426 GeV) for  $m_{G'} = 2.3 m_{\theta^0}$  ( $m_{G'} = 5 m_{\theta^0}$ ).
- <sup>4</sup> AALTONEN 13r search for new resonance decaying to  $\sigma\sigma$ , with hypothetical strongly interacting  $\sigma$  particle subsequently decaying to 2 jets, in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV, using data corresponding to an integrated luminosity of  $6.6 \text{ fb}^{-1}$ . For  $50 \text{ GeV} < m_\sigma < m_{g_A}/2$ , axigluons in mass range 150–400 GeV are excluded.
- <sup>5</sup> CHATRCHYAN 13A search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- <sup>6</sup> CHATRCHYAN 13As search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 8$  TeV.
- <sup>7</sup> CHATRCHYAN 13Au search for the pair produced color-octet vector bosons decaying to  $q\bar{q}$  pairs in  $pp$  collisions. The quoted limit is for  $B(g_A \rightarrow q\bar{q}) = 1$ .
- <sup>8</sup> ABZOV 12r search for massive color octet vector particle decaying to  $t\bar{t}$ . The quoted limit assumes  $g_A$  couplings with light quarks are suppressed by 0.2.
- <sup>9</sup> CHATRCHYAN 11v search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- <sup>10</sup> AALTONEN 10L search for massive color octet non-chiral vector particle decaying into  $t\bar{t}$  pair with mass in the range  $400 \text{ GeV} < M < 800 \text{ GeV}$ . See their Fig. 6 for limit in the mass-coupling plane.
- <sup>11</sup> KHACHATRYAN 10 search for new resonance decaying to dijets in  $pp$  collisions at  $\sqrt{s} = 7$  TeV.
- <sup>12</sup> AALTONEN 09ac search for new narrow resonance decaying to dijets.
- <sup>13</sup> CHOUDHURY 07 limit is from the  $t\bar{t}$  production cross section measured at CDF.

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

<sup>14</sup> DONCHESKI 98 compare  $\alpha_s$  derived from low-energy data and that from  $\Gamma(Z \rightarrow \text{hadrons})/\Gamma(Z \rightarrow \text{leptons})$ .  
<sup>15</sup> ABE 97G search for new particle decaying to dijets.  
<sup>16</sup> ABE 95N assume axiguons decaying to quarks in the Standard Model only.  
<sup>17</sup> ABE 93G assume  $\Gamma(g_A) = N\alpha_s m_{g_A}/6$  with  $N = 10$ .  
<sup>18</sup> CUYPERS 91 compare  $\alpha_s$  measured in  $\Upsilon$  decay and that from  $R$  at PEP/PETRA energies.  
<sup>19</sup> ABE 90H assumes  $\Gamma(g_A) = N\alpha_s m_{g_A}/6$  with  $N = 5$  ( $\Gamma(g_A) = 0.09 m_{g_A}$ ). For  $N = 10$ , the excluded region is reduced to 120–150 GeV.  
<sup>20</sup> ROBINETT 89 result demands partial-wave unitarity of  $J = 0$   $t\bar{t} \rightarrow t\bar{t}$  scattering amplitude and derives a limit  $m_{g_A} > 0.5 m_t$ . Assumes  $m_t > 56$  GeV.  
<sup>21</sup> ALBAJAR 88B result is from the nonobservation of a peak in two-jet invariant mass distribution.  $\Gamma(g_A) < 0.4 m_{g_A}$  assumed. See also BAGGER 88.  
<sup>22</sup> CUYPERS 88 requires  $\Gamma(\Upsilon \rightarrow g g_A) < \Gamma(\Upsilon \rightarrow g g g)$ . A similar result is obtained by DONCHESKI 88.  
<sup>23</sup> DONCHESKI 88B requires  $\Gamma(\Upsilon \rightarrow g q \bar{q})/\Gamma(\Upsilon \rightarrow g g g) < 0.25$ , where the former decay proceeds via axiguon exchange. A more conservative estimate of  $< 0.5$  leads to  $m_{g_A} > 21$  GeV.

### MASS LIMITS for Color-Octet Scalar Bosons

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 150–287	95	<sup>1</sup> KHACHATRYAN...15AV CMS <sup>2</sup> AAD 13K ATLS	$pp \rightarrow \Theta^0 \Theta^0 \rightarrow b\bar{b} Z g$ $pp \rightarrow S_8 S_8 X, S_8 \rightarrow 2$ jets	
<sup>1</sup> KHACHATRYAN 15AV search for pair productions of neutral color-octet weak-triplet scalar particles ( $\Theta^0$ ), decaying to $b\bar{b}$ , $Z g$ or $\gamma g$ , in $pp$ collisions at $\sqrt{s} = 8$ TeV. The $\Theta^0$ particle is often predicted in coloron ( $G'$ , color-octet gauge boson) models and appear in the $pp$ collisions through $G' \rightarrow \Theta^0 \Theta^0$ decays. Assuming $B(\Theta^0 \rightarrow b\bar{b}) = 0.5$ , they give limits $m_{\Theta^0} > 623$ GeV (426 GeV) for $m_{G'} = 2.3 m_{\Theta^0}$ ( $m_{G'} = 5 m_{\Theta^0}$ ). <sup>2</sup> AAD 13K search for pair production of color-octet scalar particles in $pp$ collisions at $\sqrt{s} = 7$ TeV. Cross section limits are interpreted as mass limits on scalar partners of a Dirac gluino.				

### X<sup>0</sup> (Heavy Boson) Searches in Z Decays

Searches for radiative transition of Z to a lighter spin-0 state  $X^0$  decaying to hadrons, a lepton pair, a photon pair, or invisible particles as shown in the comments. The limits are for the product of branching ratios.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>1</sup> BARATE 98U ALEP <sup>2</sup> ACCIARRI 97Q L3 <sup>3</sup> ACTON 93E OPAL <sup>4</sup> ABREU 92D DLPH <sup>5</sup> ADRIANI 92F L3 <sup>6</sup> ACTON 91 OPAL	$X^0 \rightarrow \ell\bar{\ell}, q\bar{q}, g g, \gamma\gamma, \nu\bar{\nu}$ $X^0 \rightarrow \text{invisible particle(s)}$ $X^0 \rightarrow \gamma\gamma$ $X^0 \rightarrow \text{hadrons}$ $X^0 \rightarrow \text{hadrons}$ $X^0 \rightarrow \text{anything}$	
$<1.1 \times 10^{-4}$	95	<sup>7</sup> ACTON 91B OPAL <sup>7</sup> ACTON 91B OPAL	$X^0 \rightarrow e^+ e^-$ $X^0 \rightarrow \mu^+ \mu^-$	
$<9 \times 10^{-5}$	95	<sup>7</sup> ACTON 91B OPAL	$X^0 \rightarrow \tau^+ \tau^-$	
$<1.1 \times 10^{-4}$	95	<sup>8</sup> ADEVA 91D L3 <sup>8</sup> ADEVA 91D L3	$X^0 \rightarrow e^+ e^-$ $X^0 \rightarrow \mu^+ \mu^-$	
$<2.8 \times 10^{-4}$	95	<sup>9</sup> ADEVA 91D L3	$X^0 \rightarrow \text{hadrons}$	
$<2.3 \times 10^{-4}$	95	<sup>10</sup> AKRAWY 90J OPAL	$X^0 \rightarrow \text{hadrons}$	
$<4.7 \times 10^{-4}$	95			
$<8 \times 10^{-4}$	95			

<sup>1</sup> BARATE 98U obtain limits on  $B(Z \rightarrow \gamma X^0)B(X^0 \rightarrow \ell\bar{\ell}, q\bar{q}, g g, \gamma\gamma, \nu\bar{\nu})$ . See their Fig. 17.  
<sup>2</sup> See Fig. 4 of ACCIARRI 97Q for the upper limit on  $B(Z \rightarrow \gamma X^0; E_{\gamma} > E_{\min})$  as a function of  $E_{\min}$ .  
<sup>3</sup> ACTON 93E give  $\sigma(e^+ e^- \rightarrow X^0 \gamma) \cdot B(X^0 \rightarrow \gamma\gamma) < 0.4$  pb (95%CL) for  $m_{X^0} = 60 \pm 2.5$  GeV. If the process occurs via s-channel  $\gamma$  exchange, the limit translates to  $\Gamma(X^0) \cdot B(X^0 \rightarrow \gamma\gamma)^2 < 20$  MeV for  $m_{X^0} = 60 \pm 1$  GeV.  
<sup>4</sup> ABREU 92D give  $\sigma_Z \cdot B(Z \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow \text{hadrons}) < (3\text{--}10)$  pb for  $m_{X^0} = 10\text{--}78$  GeV. A very similar limit is obtained for spin-1  $X^0$ .  
<sup>5</sup> ADRIANI 92F search for isolated  $\gamma$  in hadronic Z decays. The limit  $\sigma_Z \cdot B(Z \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow \text{hadrons}) < (2\text{--}10)$  pb (95%CL) is given for  $m_{X^0} = 25\text{--}85$  GeV.  
<sup>6</sup> ACTON 91 searches for  $Z \rightarrow Z^* X^0$ ,  $Z^* \rightarrow e^+ e^-, \mu^+ \mu^-, \text{ or } \nu\bar{\nu}$ . Excludes any new scalar  $X^0$  with  $m_{X^0} < 9.5$  GeV/c if it has the same coupling to  $ZZ^*$  as the MSM Higgs boson.  
<sup>7</sup> ACTON 91B limits are for  $m_{X^0} = 60\text{--}85$  GeV.  
<sup>8</sup> ADEVA 91D limits are for  $m_{X^0} = 30\text{--}89$  GeV.  
<sup>9</sup> ADEVA 91D limits are for  $m_{X^0} = 30\text{--}86$  GeV.  
<sup>10</sup> AKRAWY 90J give  $\Gamma(Z \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow \text{hadrons}) < 1.9$  MeV (95%CL) for  $m_{X^0} = 32\text{--}80$  GeV. We divide by  $\Gamma(Z) = 2.5$  GeV to get product of branching ratios. For nonresonant transitions, the limit is  $B(Z \rightarrow \gamma q \bar{q}) < 8.2$  MeV assuming three-body phase space distribution.

### MASS LIMITS for a Heavy Neutral Boson Coupling to $e^+ e^-$

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

none 55–61		<sup>1</sup> ODAKA 89	VNS	$\Gamma(X^0 \rightarrow e^+ e^-) \cdot B(X^0 \rightarrow \text{had.}) \gtrsim 0.2$ MeV
$>45$	95	<sup>2</sup> DERRICK 86	HRS	$\Gamma(X^0 \rightarrow e^+ e^-) = 6$ MeV
$>46.6$	95	<sup>3</sup> ADEVA 85	MRKJ	$\Gamma(X^0 \rightarrow e^+ e^-) = 10$ keV
$>48$	95	<sup>3</sup> ADEVA 85	MRKJ	$\Gamma(X^0 \rightarrow e^+ e^-) = 4$ MeV
		<sup>4</sup> BERGER 85B	PLUT	
none 39.8–45.5		<sup>5</sup> ADEVA 84	MRKJ	$\Gamma(X^0 \rightarrow e^+ e^-) = 10$ keV
$>47.8$	95	<sup>5</sup> ADEVA 84	MRKJ	$\Gamma(X^0 \rightarrow e^+ e^-) = 4$ MeV
none 39.8–45.2		<sup>5</sup> BEHREND 84C	CELL	
$>47$	95	<sup>5</sup> BEHREND 84C	CELL	$\Gamma(X^0 \rightarrow e^+ e^-) = 4$ MeV

<sup>1</sup> ODAKA 89 looked for a narrow or wide scalar resonance in  $e^+ e^- \rightarrow \text{hadrons}$  at  $E_{\text{cm}} = 55.0\text{--}60.8$  GeV.  
<sup>2</sup> DERRICK 86 found no deviation from the Standard Model Bhabha scattering at  $E_{\text{cm}} = 29$  GeV and set limits on the possible scalar boson  $e^+ e^-$  coupling. See their figure 4 for excluded region in the  $\Gamma(X^0 \rightarrow e^+ e^-) - m_{X^0}$  plane. Electronic chiral invariance requires a parity doublet of  $X^0$ , in which case the limit applies for  $\Gamma(X^0 \rightarrow e^+ e^-) = 3$  MeV.  
<sup>3</sup> ADEVA 85 first limit is from  $2\gamma, \mu^+ \mu^-, \text{ hadrons}$  assuming  $X^0$  is a scalar. Second limit is from  $e^+ e^-$  channel.  $E_{\text{cm}} = 40\text{--}47$  GeV. Supersedes ADEVA 84.  
<sup>4</sup> BERGER 85B looked for effect of spin-0 boson exchange in  $e^+ e^- \rightarrow e^+ e^-$  and  $\mu^+ \mu^-$  at  $E_{\text{cm}} = 34.7$  GeV. See Fig. 5 for excluded region in the  $m_{X^0} - \Gamma(X^0)$  plane.  
<sup>5</sup> ADEVA 84 and BEHREND 84C have  $E_{\text{cm}} = 39.8\text{--}45.5$  GeV. MARK-J searched  $X^0$  in  $e^+ e^- \rightarrow \text{hadrons}$ ,  $2\gamma, \mu^+ \mu^-, e^+ e^-$  and CELLO in the same channels plus  $\tau$  pair. No narrow or broad  $X^0$  is found in the energy range. They also searched for the effect of  $X^0$  with  $m_X > E_{\text{cm}}$ . The second limits are from Bhabha data and for spin-0 singlet. The same limits apply for  $\Gamma(X^0 \rightarrow e^+ e^-) = 2$  MeV if  $X^0$  is a spin-0 doublet. The second limit of BEHREND 84C was read off from their figure 2. The original papers also list limits in other channels.

### Search for X<sup>0</sup> Resonance in $e^+ e^-$ Collisions

The limit is for  $\Gamma(X^0 \rightarrow e^+ e^-) \cdot B(X^0 \rightarrow f)$ , where  $f$  is the specified final state. Spin 0 is assumed for  $X^0$ .

VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<10^3$	95	<sup>1</sup> ABE 93C	VNS	$\Gamma(e e)$
$<(0.4\text{--}10)$	95	<sup>2</sup> ABE 93C	VNS	$f = \gamma\gamma$
$<(0.3\text{--}5)$	95	<sup>3,4</sup> ABE 93D	TOPZ	$f = \gamma\gamma$
$<(2\text{--}12)$	95	<sup>3,4</sup> ABE 93D	TOPZ	$f = \text{hadrons}$
$<(4\text{--}200)$	95	<sup>4,5</sup> ABE 93D	TOPZ	$f = e e$
$<(0.1\text{--}6)$	95	<sup>4,5</sup> ABE 93D	TOPZ	$f = \mu\mu$
$<(0.5\text{--}8)$	90	<sup>6</sup> STERNER 93	AMY	$f = \gamma\gamma$

<sup>1</sup> Limit is for  $\Gamma(X^0 \rightarrow e^+ e^-) m_{X^0} = 56\text{--}63.5$  GeV for  $\Gamma(X^0) = 0.5$  GeV.  
<sup>2</sup> Limit is for  $m_{X^0} = 56\text{--}61.5$  GeV and is valid for  $\Gamma(X^0) \ll 100$  MeV. See their Fig. 5 for limits for  $\Gamma = 1, 2$  GeV.  
<sup>3</sup> Limit is for  $m_{X^0} = 57.2\text{--}60$  GeV.  
<sup>4</sup> Limit is valid for  $\Gamma(X^0) \ll 100$  MeV. See paper for limits for  $\Gamma = 1$  GeV and those for  $J = 2$  resonances.  
<sup>5</sup> Limit is for  $m_{X^0} = 56.6\text{--}60$  GeV.  
<sup>6</sup> STERNER 93 limit is for  $m_{X^0} = 57\text{--}59.6$  GeV and is valid for  $\Gamma(X^0) < 100$  MeV. See their Fig. 2 for limits for  $\Gamma = 1, 3$  GeV.

### Search for X<sup>0</sup> Resonance in $e p$ Collisions

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	<sup>1</sup> CHEKANOV 02B	ZEUS	$X \rightarrow jj$

<sup>1</sup> CHEKANOV 02B search for photoproduction of X decaying into dijets in  $e p$  collisions. See their Fig. 5 for the limit on the photoproduction cross section.

### Search for X<sup>0</sup> Resonance in $e^+ e^- \rightarrow X^0 \gamma$

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	<sup>1</sup> ABBIENDI 03D	OPAL	$X^0 \rightarrow \gamma\gamma$
	<sup>2</sup> ABREU 00Z	DLPH	$X^0$ decaying invisibly
	<sup>3</sup> ADAM 96C	DLPH	$X^0$ decaying invisibly

<sup>1</sup> ABBIENDI 03D measure the  $e^+ e^- \rightarrow \gamma\gamma\gamma$  cross section at  $\sqrt{s} = 181\text{--}209$  GeV. The upper bound on the production cross section,  $\sigma(e^+ e^- \rightarrow X^0 \gamma)$  times the branching ratio for  $X^0 \rightarrow \gamma\gamma$ , is less than 0.03 pb at 95%CL for  $X^0$  masses between 20 and 180 GeV. See their Fig. 9b for the limits in the mass-cross section plane.  
<sup>2</sup> ABREU 00Z is from the single photon cross section at  $\sqrt{s} = 183, 189$  GeV. The production cross section upper limit is less than 0.3 pb for  $X^0$  mass between 40 and 160 GeV. See their Fig. 4 for the limit in mass-cross section plane.  
<sup>3</sup> ADAM 96C is from the single photon production cross at  $\sqrt{s} = 130, 136$  GeV. The upper bound is less than 3 pb for  $X^0$  masses between 60 and 130 GeV. See their Fig. 5 for the exact bound on the cross section  $\sigma(e^+ e^- \rightarrow \gamma X^0)$ .

### Search for X<sup>0</sup> Resonance in $Z \rightarrow f \bar{f} X^0$

The limit is for  $B(Z \rightarrow f \bar{f} X^0) \cdot B(X^0 \rightarrow F)$  where  $f$  is a fermion and  $F$  is the specified final state. Spin 0 is assumed for  $X^0$ .

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

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# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons

$<3.7 \times 10^{-6}$	95	1 ABREU 96T DLPH $f=e, \mu, \tau; F=\gamma\gamma$
		2 ABREU 96T DLPH $f=\nu; F=\gamma\gamma$
		3 ABREU 96T DLPH $f=q; F=\gamma\gamma$
$<6.8 \times 10^{-6}$	95	2 ACTON 93E OPAL $f=e, \mu, \tau; F=\gamma\gamma$
$<5.5 \times 10^{-6}$	95	2 ACTON 93E OPAL $f=q; F=\gamma\gamma$
$<3.1 \times 10^{-6}$	95	2 ACTON 93E OPAL $f=\nu; F=\gamma\gamma$
$<6.5 \times 10^{-6}$	95	2 ACTON 93E OPAL $f=e, \mu; F=\ell\bar{\ell}, q\bar{q}, \nu\bar{\nu}$
$<7.1 \times 10^{-6}$	95	2 BUSKULIC 93F ALEP $f=e, \mu; F=\ell\bar{\ell}, q\bar{q}, \nu\bar{\nu}$
		4 ADRIANI 92F L3 $f=q; F=\gamma\gamma$

1 ABREU 96T obtain limit as a function of  $m_{X^0}$ . See their Fig. 6.  
2 Limit is for  $m_{X^0}$  around 60 GeV.  
3 ABREU 96T obtain limit as a function of  $m_{X^0}$ . See their Fig. 15.  
4 ADRIANI 92F give  $\sigma_{\gamma} \cdot B(Z \rightarrow q\bar{q}X^0) \cdot B(X^0 \rightarrow \gamma\gamma) < (0.75-1.5) \text{ pb (95\%CL) for } m_{X^0} = 10-70 \text{ GeV. The limit is 1 pb at 60 GeV.}$

### Search for $X^0$ Resonance in $WX^0$ final state

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	1 AALTONEN 13AA CDF $X^0 \rightarrow jj$		
	2 CHATRCHYAN 12BR CMS $X^0 \rightarrow jj$		
	3 ABABOV 11I D0 $X^0 \rightarrow jj$		
	4 ABE 97W CDF $X^0 \rightarrow b\bar{b}$		
1 AALTONEN 13AA search for $X^0$ production associated with $W$ (or $Z$ ) in $p\bar{p}$ collisions at $E_{\text{cm}} = 1.96 \text{ TeV}$ . The upper limit on the cross section $\sigma(p\bar{p} \rightarrow WX^0)$ is 2.2 pb for $M_{X^0} = 145 \text{ GeV}$ .			
2 CHATRCHYAN 12BR search for $X^0$ production associated with $W$ in $pp$ collisions at $E_{\text{cm}} = 7 \text{ TeV}$ . The upper limit on the cross section is 5.0 pb at 95% CL for $m_{X^0} = 150 \text{ GeV}$ .			
3 ABABOV 11I search for $X^0$ production associated with $W$ in $p\bar{p}$ collisions at $E_{\text{cm}} = 1.96 \text{ TeV}$ . The 95% CL upper limit on the cross section ranges from 2.57 to 1.28 pb for $X^0$ mass between 110 and 170 GeV.			
4 ABE 97W search for $X^0$ production associated with $W$ in $p\bar{p}$ collisions at $E_{\text{cm}} = 1.8 \text{ TeV}$ . The 95%CL upper limit on the production cross section times the branching ratio for $X^0 \rightarrow b\bar{b}$ ranges from 14 to 19 pb for $X^0$ mass between 70 and 120 GeV. See their Fig. 3 for upper limits of the production cross section as a function of $m_{X^0}$ .			

### Search for $X^0$ Resonance in Quarkonium Decays

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3 \times 10^{-5} - 6 \times 10^{-3}$	90	1 BALEST 95 CLE2 $\Upsilon(1S) \rightarrow X^0 \bar{X}^0 \gamma$ $m_{X^0} < 3.9 \text{ GeV}$		
1 BALEST 95 three-body limit is for phase-space photon energy distribution and angular distribution same as for $T \rightarrow gg\gamma$ .				

### REFERENCES FOR Searches for New Heavy Bosons ( $W'$ , $Z'$ , leptoquarks, etc.)

AAD 16G EPJ C75 6	G. Aad et al.	(ATLAS Collab.)
KHACHATRY... 16G PR D93 012001	V. Khachatryan et al.	(CMS Collab.)
AAD 15AM JHEP 1507 157	G. Aad et al.	(CMS Collab.)
AAD 15AO JHEP 1508 148	G. Aad et al.	(ATLAS Collab.)
AAD 15AT EPJ C75 79	G. Aad et al.	(ATLAS Collab.)
AAD 15AU EPJ C75 69	G. Aad et al.	(ATLAS Collab.)
AAD 15AV EPJ C75 165	G. Aad et al.	(ATLAS Collab.)
AAD 15AZ EPJ C75 209	G. Aad et al.	(ATLAS Collab.)
AAD 15BB EPJ C75 263	G. Aad et al.	(ATLAS Collab.)
AAD 15CD PR D92 092001	G. Aad et al.	(ATLAS Collab.)
AAD 15CP JHEP 1512 055	G. Aad et al.	(ATLAS Collab.)
AAD 15O PRL 115 031801	G. Aad et al.	(ATLAS Collab.)
AAD 15R PL B743 235	G. Aad et al.	(ATLAS Collab.)
AAD 15V PR D91 052007	G. Aad et al.	(ATLAS Collab.)
AALTONEN 15C PRL 115 061801	T. Aaltonen et al.	(CDF Collab.)
BESSAA 15 EPJ C75 97	A. Bessaa, S. Davidson	(CMS Collab.)
KHACHATRY... 15AE JHEP 1504 025	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15AJ JHEP 1507 042	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15AV JHEP 1509 201	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15C PL B740 83	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15F PRL 114 101801	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15O PL B748 255	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15T PR D91 092005	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 15V PR D91 052009	V. Khachatryan et al.	(CMS Collab.)
SAHOO 15A PR D91 094019	S. Sahoo, R. Mohanta	(ATLAS Collab.)
AAD 14AI JHEP 1409 037	G. Aad et al.	(ATLAS Collab.)
AAD 14AT PL B738 428	G. Aad et al.	(ATLAS Collab.)
AAD 14S PL B737 223	G. Aad et al.	(ATLAS Collab.)
AAD 14V PR D90 052005	G. Aad et al.	(ATLAS Collab.)
KHACHATRY... 14 JHEP 1408 173	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 14A JHEP 1408 174	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 14O EPJ C74 3149	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 14T PL B739 229	V. Khachatryan et al.	(CMS Collab.)
MARTINEZ 14 PR D90 015028	R. Martinez, F. Ochoa	(LOUV, ETH, PSI+)
PRIEELS 14 PR D90 112003	R. Priels et al.	(ATLAS Collab.)
AAD 13AE JHEP 1306 033	G. Aad et al.	(ATLAS Collab.)
AAD 13AI PL B723 15	G. Aad et al.	(ATLAS Collab.)
AAD 13AO PR D87 112006	G. Aad et al.	(ATLAS Collab.)
AAD 13AQ PR D88 012004	G. Aad et al.	(ATLAS Collab.)
AAD 13D JHEP 1301 029	G. Aad et al.	(ATLAS Collab.)
AAD 13G JHEP 1301 116	G. Aad et al.	(ATLAS Collab.)
AAD 13K EPJ C73 2263	G. Aad et al.	(ATLAS Collab.)
AAD 13S PL B719 242	G. Aad et al.	(ATLAS Collab.)
AALTONEN 13A PRL 110 121802	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 13AA PR D88 092004	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 13R PRL 111 031802	T. Aaltonen et al.	(CDF Collab.)
CHATRCHYAN 13A JHEP 1301 013	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AF PL B720 63	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AJ PL B723 280	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AP PR D87 072002	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AQ PR D87 072005	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AS PR D87 114015	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AU PRL 110 141802	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13BM PRL 111 211804	S. Chatrchyan et al.	(CMS Collab.)
Also PRL 112 119903 (err.)	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13E PL B718 1229	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13M PRL 110 081801	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13U JHEP 1302 036	S. Chatrchyan et al.	(CMS Collab.)
SAKAKI 13 PR D88 094012	Y. Sakaki et al.	(ATLAS Collab.)
AAD 12AV PRL 109 081801	G. Aad et al.	(ATLAS Collab.)
AAD 12BB PR D85 112012	G. Aad et al.	(ATLAS Collab.)
AAD 12BV JHEP 1209 041	G. Aad et al.	(ATLAS Collab.)
AAD 12CC JHEP 1211 138	G. Aad et al.	(ATLAS Collab.)
AAD 12CK PR D86 091103	G. Aad et al.	(ATLAS Collab.)
AAD 12CR EPJ C72 2241	G. Aad et al.	(ATLAS Collab.)
AAD 12H PL B709 158	G. Aad et al.	(ATLAS Collab.)
Also PL B711 442 (err.)	G. Aad et al.	(ATLAS Collab.)
AAD 12K EPJ C72 2083	G. Aad et al.	(ATLAS Collab.)
AAD 12M EPJ C72 2056	G. Aad et al.	(ATLAS Collab.)
AAD 12O EPJ C72 2151	G. Aad et al.	(ATLAS Collab.)
AALTONEN 12AR PR D86 112002	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 12N PRL 108 211805	T. Aaltonen et al.	(CDF Collab.)
ABABOV 12R PR D85 051101	V.M. Abazov et al.	(DO Collab.)
ABRAMOWICZ 12A PR D86 012005	H. Abramowicz et al.	(ZEUS Collab.)
CHATRCHYAN 12AF PRL 109 141801	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12AG PR D86 052013	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12AI JHEP 1208 110	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12AQ JHEP 1209 029	S. Chatrchyan et al.	(CMS Collab.)
Also JHEP 1403 132 (err.)	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12AR PL B717 351	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12BG PRL 109 261802	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12BL JHEP 1212 015	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12BO JHEP 1212 055	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12BR PRL 109 251801	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12M PL B714 158	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12O PL B716 82	S. Chatrchyan et al.	(CMS Collab.)
KOSNIK 12 PR D86 055004	N. Kosnik	(LALO, STFN)
AAD 11D PR D83 112006	G. Aad et al.	(ATLAS Collab.)
AAD 11H PRL 106 251801	G. Aad et al.	(ATLAS Collab.)
AAD 11Z EPJ C71 1809	G. Aad et al.	(ATLAS Collab.)
AALTONEN 11AD PR D84 072003	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 11AE PR D84 072004	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 11C PR D83 031102	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 11I PRL 106 121801	T. Aaltonen et al.	(CDF Collab.)
AARON 11A PL B701 20	F. D. Aaron et al.	(H1 Collab.)
AARON 11B PL B704 388	F. D. Aaron et al.	(H1 Collab.)
AARON 11C PL B705 52	F. D. Aaron et al.	(H1 Collab.)
ABABOV 11A PL B695 88	V.M. Abazov et al.	(DO Collab.)
ABABOV 11H PRL 107 011801	V. M. Abazov et al.	(DO Collab.)
ABABOV 11I PRL 107 011804	V. M. Abazov et al.	(DO Collab.)
ABABOV 11L PL B699 145	V. M. Abazov et al.	(DO Collab.)
ABABOV 11V PR D84 071104	V. M. Abazov et al.	(DO Collab.)
BUENO 11 PR D84 032005	J.F. Bueno et al.	(TWIST Collab.)
Also PR D85 039908 (err.)	J.F. Bueno et al.	(TWIST Collab.)
CHATRCHYAN 11N PL B703 246	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 11O JHEP 1108 005	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 11Y PL B704 123	S. Chatrchyan et al.	(CMS Collab.)
DORSNER 11 JHEP 1111 002	I. Dorsner et al.	(CMS Collab.)
KHACHATRY... 11D PRL 106 201802	V. Khachatryan et al.	(CMS Collab.)
KHACHATRY... 11E PRL 106 201803	V. Khachatryan et al.	(CMS Collab.)
AALTONEN 10L PL B691 183	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 10N PRL 104 241801	T. Aaltonen et al.	(CDF Collab.)
ABABOV 10L PL B691 185	V.M. Abazov et al.	(DO Collab.)
DEL-AGUILA 10 JHEP 1009 033	F. del Aguila, M. de Blas, M. Perez-Victoria	(GRAN)
KHACHATRY... 10 PRL 105 211801	V. Khachatryan et al.	(CMS Collab.)
Also PRL 106 029902	V. Khachatryan et al.	(CMS Collab.)
WAUTERS 10 PR C82 055502	F. Wauters et al.	(REZ, TAMU)
AALTONEN 09AC PR D79 112002	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 09T PRL 102 031801	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 09V PRL 102 091805	T. Aaltonen et al.	(CDF Collab.)
ABABOV 09 PL B671 224	V.M. Abazov et al.	(DO Collab.)
ABABOV 09AF PL B681 224	V.M. Abazov et al.	(DO Collab.)
ERLER 09 JHEP 0908 017	J. Erler et al.	(CDF Collab.)
AALTONEN 08D PR D77 051102	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 08P PR D77 091105	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 08Y PRL 100 231801	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 08Z PRL 101 071802	T. Aaltonen et al.	(CDF Collab.)
ABABOV 08AA PL B668 98	V.M. Abazov et al.	(DO Collab.)
ABABOV 08AD PL B668 357	V.M. Abazov et al.	(DO Collab.)
ABABOV 08AN PRL 101 241802	V.M. Abazov et al.	(DO Collab.)
ABABOV 08C PRL 100 031804	V.M. Abazov et al.	(DO Collab.)
MACDONALD 08 PR D78 032010	R.P. MacDonald et al.	(TWIST Collab.)
ZHANG 08 NP B802 247	Y. Zhang et al.	(PKGU, UMD)
AALTONEN 07H PRL 99 171802	T. Aaltonen et al.	(CDF Collab.)
ABABOV 07E PL B647 74	V.M. Abazov et al.	(DO Collab.)
ABABOV 07J PRL 99 061801	V.M. Abazov et al.	(DO Collab.)
AKTAS 07A EPJ C52 833	A. Aktas et al.	(H1 Collab.)
CHOUDHURY 07 PL B657 69	D. Choudhury et al.	(TRIMF)
MELCONIAN 07 PL B649 370	D. Melconian et al.	(ALPH Collab.)
SCHAEF 07A EPJ C49 411	S. Schaefer et al.	(HEID, ILLG, KARL+)
SCHUMANN 07 PRL 99 191803	M. Schumann et al.	(CDF Collab.)
SMIRNOV 07 MPL A22 2353	A.D. Smirnov	(CDF Collab.)
ABABOV 06A PL B636 183	V.M. Abazov et al.	(DO Collab.)
ABABOV 06L PL B640 230	V.M. Abazov et al.	(DO Collab.)
ABDALLAH 06C EPJ C45 589	J. Abdallah et al.	(DELPHI Collab.)
ABULENCIA 06A PRL 96 211801	A. Abulencia et al.	(CDF Collab.)
ABULENCIA 06B PRL 96 211802	A. Abulencia et al.	(CDF Collab.)
ABULENCIA 06C PR D73 051102	A. Abulencia et al.	(CDF Collab.)
ABABOV 05H PR D71 071104	V.M. Abazov et al.	(DO Collab.)
ABULENCIA 05A PRL 95 252001	A. Abulencia et al.	(CDF Collab.)
ACOSTA 05I PR D71 112001	D. Acosta et al.	(CDF Collab.)
ACOSTA 05P PR D72 051107	D. Acosta et al.	(CDF Collab.)
ACOSTA 05R PRL 95 131801	D. Acosta et al.	(CDF Collab.)
AKTAS 05B PL B629 9	A. Aktas et al.	(H1 Collab.)
CEKANOV 05 PL B610 212	S. Cekanov et al.	(HERA ZEUS Collab.)
CEKANOV 05A EPJ C44 463	S. Cekanov et al.	(ZEUS Collab.)
CYBURT 05 ASP 23 313	R.H. Cyburt et al.	(DO Collab.)
ABABOV 04A PRL 92 221801	V.M. Abazov et al.	(DO Collab.)
ABABOV 04C PR D69 111101	V.M. Abazov et al.	(DO Collab.)
ABBIENDI 04G EPJ C33 173	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI 03D EPJ C26 331	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI 03R EPJ C31 281	G. Abbiendi et al.	(OPAL Collab.)
ACOSTA 03B PRL 90 081802	D. Acosta et al.	(CDF Collab.)
ADLOFF 03 PL B568 35	C. Adloff et al.	(H1 Collab.)
BARGER 03B PR D67 075009	V. Barger, P. Langacker, H. Lee	(BELLE Collab.)
CHANG 03 PR D68 111101	M.-C. Chang et al.	(ZEUS Collab.)
CEKANOV 03B PR D68 052004	S. Cekanov et al.	(ZEUS Collab.)
ABABOV 02B PRL 88 191801	V.M. Abazov et al.	(DO Collab.)
ABBIENDI 02B PL B526 233	G. Abbiendi et al.	(OPAL Collab.)
AFFOLDER 02C PRL 88 071806	T. Affolder et al.	(CDF Collab.)

# Gauge & Higgs Boson Particle Listings

## New Heavy Bosons, Axions ( $A^0$ ) and Other Very Light Bosons

CHEKANOV	02	PR D65 092004	S. Chekanov <i>et al.</i>	(ZEUS Collab.)	RIZZO	91	PR D44 202	T.G. Rizzo	(WISC, ISU)
CHEKANOV	02B	PL B531 9	S. Chekanov <i>et al.</i>	(ZEUS Collab.)	WALKER	91	APJ 376 51	T.P. Walker <i>et al.</i>	(HSCA, OSU, CHIC+)
MUECK	02	PR D65 085037	A. Mueck, A. Pilaftsis, R. Rueckl		ABE	90F	PL B246 297	K. Abe <i>et al.</i>	(VENUS Collab.)
ABAZOV	01B	PRL 87 061802	V.M. Abazov <i>et al.</i>	(DO Collab.)	ABE	90H	PR D41 1722	F. Abe <i>et al.</i>	(CDF Collab.)
ABAZOV	01D	PR D64 092004	V.M. Abazov <i>et al.</i>	(DO Collab.)	AKRAWY	90J	PL B246 285	M.Z. Akrawy <i>et al.</i>	(OPAL Collab.)
ADLOFF	01C	PL B523 234	C. Adloff <i>et al.</i>	(H1 Collab.)	GONZALEZ-G.	90D	PL B240 163	M.C. Gonzalez-Garcia, J.W.F. Valle	(VALE)
AFFOLDER	01I	PRL 87 231803	T. Affolder <i>et al.</i>	(CDF Collab.)	GRIFOLS	90	NP B331 244	J.A. Grifols, E. Masso	(BARC)
BREITWEG	01	PR D63 052002	J. Breitweg <i>et al.</i>	(ZEUS Collab.)	GRIFOLS	90D	PR D42 3293	J.A. Grifols, E. Masso, T.G. Rizzo	(BARC, CERN+)
CHEUNG	01B	PL B517 167	K. Cheung		KIM	90	PL B240 243	G.N. Kim <i>et al.</i>	(AMY Collab.)
THOMAS	01	NP A594 559	E. Thomas <i>et al.</i>		LOPEZ	90	PL B241 392	J.L. Lopez, D.V. Nanopoulos	(TAMU)
ABBIENDI	00M	EPJ C13 15	G. Abbiendi <i>et al.</i>	(OPAL Collab.)	BARBIERI	89B	PR D39 1229	R. Barbieri, R.N. Mohapatra	(PISA, UMD)
ABBOTT	00C	PRL 84 2088	B. Abbott <i>et al.</i>	(DO Collab.)	LANGACKER	89B	PR D40 1569	P. Langacker, S. Uma Sankar	(PENN)
ABE	00	PRL 84 5716	F. Abe <i>et al.</i>	(CDF Collab.)	ODAKA	89	JPSJ 58 3037	S. Odaka <i>et al.</i>	(VENUS Collab.)
ABREU	00S	PL B485 45	P. Abreu <i>et al.</i>	(DELPHI Collab.)	ROBINETT	89	PR D39 834	R.W. Robinett	(PSU)
ABREU	00Z	EPJ C17 53	P. Abreu <i>et al.</i>	(DELPHI Collab.)	ALBAJAR	88B	PL B209 127	C. Albajar <i>et al.</i>	(UA1 Collab.)
ACCIARRI	00P	PL B489 81	M. Acciarri <i>et al.</i>	(L3 Collab.)	BAGGER	88	PR D37 1188	J. Bagger, C. Schmidt, S. King	(HARV, BOST)
ADLOFF	00	PL B479 358	C. Adloff <i>et al.</i>	(H1 Collab.)	BALKE	88	PR D37 587	B. Balke <i>et al.</i>	(LBL, UCB, COLO, NWES+)
AFFOLDER	00K	PRL 85 2056	T. Affolder <i>et al.</i>	(CDF Collab.)	BERGSTROM	88	PL B212 386	L. Bergstrom	(STOH)
BARATE	00I	EPJ C12 183	R. Barate <i>et al.</i>	(ALEPH Collab.)	CUYPERS	88	PRL 60 1237	F. Cuyper, P.H. Frampton	(UNCCH)
BARGER	00	PL B480 149	V. Barger, K. Cheung		DONCHESKI	88	PL B206 137	M.A. Doncheski, H. Grotch, R. Robinett	(PSU)
BREITWEG	00E	EPJ C16 253	J. Breitweg <i>et al.</i>	(ZEUS Collab.)	DONCHESKI	88B	PR D38 412	M.A. Doncheski, H. Grotch, R.W. Robinett	(PSU)
CHAY	00	PR D61 035002	J. Chay, K.Y. Lee, S. Nam		BARTEL	87B	ZPHY C36 15	W. Bartel <i>et al.</i>	(JADE Collab.)
CHO	00	MPL A15 311	G. Cho		BEHREND	85B	PL B178 452	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
CORNET	00	PR D61 037701	F. Cornet, M. Relano, J. Rico		DERRICK	86	PL B166 463	M. Derrick <i>et al.</i>	(HRS Collab.)
DELGADO	00	JHEP 0001 030	A. Delgado, A. Pomarol, M. Quiros		Also		PR D34 3286	M. Derrick <i>et al.</i>	(HRS Collab.)
ERLER	00	PRL 84 212	J. Erler, P. Langacker		JODIDIO	86	PR D34 1967	A. Jodidio <i>et al.</i>	(LBL, NWES, TRIU)
GABRIELLI	00	PR D62 055009	E. Gabrielli		Also		PR D37 237 (erratum)	A. Jodidio <i>et al.</i>	(LBL, NWES, TRIU)
RIZZO	00	PR D61 016007	T.G. Rizzo, J.D. Wells		MOHAPATRA	86	PR D34 909	R.N. Mohapatra	(UMD)
ROSNER	00	PR D61 016006	J.L. Rosner		ADEVA	85	PL B52B 439	B. Adeva <i>et al.</i>	(Mark-I Collab.)
ZARNECKI	00	EPJ C17 695	A. Zarnecki		BERGER	85B	ZPHY C27 341	C. Berger <i>et al.</i>	(PLUTO Collab.)
ABBIENDI	99	EPJ C6 1	G. Abbiendi <i>et al.</i>	(OPAL Collab.)	STOKER	85B	PRL 54 1887	D.P. Stoker <i>et al.</i>	(LBL, NWES, TRIU)
ABBOTT	99J	PRL 83 2836	B. Abbott <i>et al.</i>	(DO Collab.)	ADDA	84	PRL 53 134	B. Adeva <i>et al.</i>	(Mark-I Collab.)
ABREU	99G	PL B446 62	P. Abreu <i>et al.</i>	(DELPHI Collab.)	BEHREND	84C	PL B140B 130	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
ACKERSTAFF	99D	EPJ C8 3	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)	BERGSMAS	83	PL B122B 465	F. Bergsma <i>et al.</i>	(CHARM Collab.)
ADLOFF	99	EPJ C11 447	C. Adloff <i>et al.</i>	(H1 Collab.)	CARR	83	PRL 51 627	J. Carr <i>et al.</i>	(LBL, NWES, TRIU)
Also		EPJ C14 553 (errat.)	C. Adloff <i>et al.</i>	(H1 Collab.)	BEALL	82	PRL 48 848	G. Beall, M. Bander, A. Soni	(UCI, UCLA)
CASALBUONI	99	PL B460 135	R. Casalbuoni <i>et al.</i>		SHANKER	82	NP B204 375	O. Shanker	(TRIUMF)
CZAKON	99	PL B458 355	M. Czakon, J. Gluza, M. Zralek						
ERLER	99	PL B456 68	J. Erler, P. Langacker						
MARCIANO	99	PR D60 093006	W. Marciano						
MASIP	99	PR D60 096005	M. Masip, A. Pomarol						
NATH	99	PR D60 116004	P. Nath, M. Yamaguchi						
STRUMIA	99	PL B466 107	A. Strumia						
ABBOTT	98E	PL B466 107	B. Abbott <i>et al.</i>	(DO Collab.)					
ABBOTT	98J	PRL 81 38	B. Abbott <i>et al.</i>	(DO Collab.)					
ABE	98S	PRL 81 4806	F. Abe <i>et al.</i>	(CDF Collab.)					
ABE	98V	PRL 81 5742	F. Abe <i>et al.</i>	(CDF Collab.)					
ACCIARRI	98J	PL B433 163	M. Acciarri <i>et al.</i>	(L3 Collab.)					
ACKERSTAFF	98V	EPJ C2 441	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)					
BARATE	98U	EPJ C4 571	R. Barate <i>et al.</i>	(ALEPH Collab.)					
BARENBOIM	98	EPJ C1 369	G. Baranboim						
CHO	98	EPJ C5 155	G. Cho, K. Hagiwara, S. Matsumoto						
CONRAD	98	RMP 70 1341	J.M. Conrad, M.H. Shaevitz, T. Bolton						
DONCHESKI	98	PR D58 097702	M.A. Doncheski, R.W. Robinett						
GROSS-PILCHER	98	hep-ex/9810015	C. Grosso-Pilcher, G. Landsberg, M. Paterno						
ABE	97F	PRL 78 2906	F. Abe <i>et al.</i>	(CDF Collab.)					
ABE	97G	PR D55 R5263	F. Abe <i>et al.</i>	(CDF Collab.)					
ABE	97S	PRL 79 2192	F. Abe <i>et al.</i>	(CDF Collab.)					
ABE	97W	PRL 79 3819	F. Abe <i>et al.</i>	(CDF Collab.)					
ABE	97X	PRL 79 4327	F. Abe <i>et al.</i>	(CDF Collab.)					
ACCIARRI	97Q	PL B412 201	M. Acciarri <i>et al.</i>	(L3 Collab.)					
ARIMA	97	PR D55 19	T. Arima <i>et al.</i>	(VENUS Collab.)					
BARENBOIM	97	PR D55 4213	G. Baranboim <i>et al.</i>	(VALE, IFIC)					
DEANDREA	97	PL B409 277	A. Deandrea	(MARS)					
DERRICK	97	ZPHY C73 613	M. Derrick <i>et al.</i>	(ZEUS Collab.)					
Y. GROSSMAN	97	PR D55 2765	Y. Grossman, Z. Ligeti, E. Nardi	(REHO, CIT)					
JADACH	97	PL B408 281	S. Jadach, B.F.L. Ward, Z. Was	(CERN, IMPK+)					
STAHL	97	ZPHY C74 73	A. Stahl, H. Voss	(BONN)					
ABACHI	96C	PRL 76 3271	S. Abachi <i>et al.</i>	(DO Collab.)					
ABREU	96T	ZPHY C72 179	P. Abreu <i>et al.</i>	(DELPHI Collab.)					
ADAM	96C	PL B380 471	W. Adam <i>et al.</i>	(DELPHI Collab.)					
AID	96B	PL B369 173	S. Aid <i>et al.</i>	(H1 Collab.)					
ALLET	96	PL B383 139	M. Allet <i>et al.</i>	(VILL, LEUV, LOUV, WISC)					
ABACHI	95E	PL B358 405	S. Abachi <i>et al.</i>	(DO Collab.)					
ABE	95N	PRL 74 3538	F. Abe <i>et al.</i>	(CDF Collab.)					
BALEST	95	PR D51 2053	F. Balest <i>et al.</i>	(CLEO Collab.)					
KUZNETSOV	95	PRL 75 794	I.A. Kuznetsov <i>et al.</i>	(PNPI, KIAE, HARV+)					
KUZNETSOV	95B	PAN 58 2113	A.V. Kuznetsov, N.V. Mikheev	(YARO)					
Translated from YAF	58	2228							
MIZUKOSHI	95	NP B443 20	J.K. Mizukoshi, O.J.P. Eboli, M.C. Gonzalez-Garcia						
ABREU	94O	ZPHY C64 183	P. Abreu <i>et al.</i>	(DELPHI Collab.)					
BHATTACH...	94	PL B336 100	G. Bhattacharyya, J. Ellis, K. Sridhar	(CERN)					
Also		PL B338 522 (erratum)	G. Bhattacharyya, J. Ellis, K. Sridhar	(CERN)					
BHATTACH...	94B	PL B338 522 (erratum)	G. Bhattacharyya, J. Ellis, K. Sridhar	(CERN)					
DAVIDSON	94	ZPHY C61 613	S. Davidson, D. Bailey, B.A. Campbell	(CFPA+)					
KUZNETSOV	94	PL B329 295	A.V. Kuznetsov, N.V. Mikheev	(YARO)					
KUZNETSOV	94B	JETPL 60 315	I.A. Kuznetsov <i>et al.</i>	(PNPI, KIAE, HARV+)					
Translated from ZETFP	60	311							
LEURER	94	PR D50 536	M. Leuner	(REHO)					
LEURER	94B	PR D49 333	M. Leuner	(REHO)					
Also		PRL 71 1324	M. Leuner	(REHO)					
MAHANTA	94	PL B337 128	U. Mahanta	(MEHTA)					
SEVERIJNS	94	PRL 73 611 (erratum)	N. Severijns <i>et al.</i>	(LOUV, WISC, LEUV+)					
VILAIN	94B	PL B332 465	P. Vilain <i>et al.</i>	(CHARM II Collab.)					
ABE	93C	PL B302 119	K. Abe <i>et al.</i>	(VENUS Collab.)					
T. ABE	93D	PL B304 373	T. Abe <i>et al.</i>	(TOPAZ Collab.)					
ABE	93G	PRL 71 2542	F. Abe <i>et al.</i>	(CDF Collab.)					
ABREU	93J	PL B316 620	P. Abreu <i>et al.</i>	(DELPHI Collab.)					
ACTON	93E	PL B311 391	P.D. Acton <i>et al.</i>	(OPAL Collab.)					
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)					
ALITTI	93	NP B400 3	J. Alitti <i>et al.</i>	(UA2 Collab.)					
BHATTACH...	93	PR D47 R3693	G. Bhattacharyya <i>et al.</i>	(CALC, JADA, ICTP+)					
BUSKULIC	93F	PL B308 425	D. Buskulic <i>et al.</i>	(ALEPH Collab.)					
DERRICK	93	PL B306 173	M. Derrick <i>et al.</i>	(ZEUS Collab.)					
RIZZO	93	PR D48 4470	T.G. Rizzo	(ANL)					
SEVERIJNS	93	PRL 70 4047	N. Severijns <i>et al.</i>	(LOUV, WISC, LEUV+)					
Also		PRL 73 611 (erratum)	N. Severijns <i>et al.</i>	(LOUV, WISC, LEUV+)					
STERNER	93	PL B303 385	K.L. Sterner <i>et al.</i>	(AMY Collab.)					
ABREU	92D	ZPHY C53 555	P. Abreu <i>et al.</i>	(DELPHI Collab.)					
ADRIANI	92F	PL B292 472	O. Adriani <i>et al.</i>	(L3 Collab.)					
DECAMP	92	PRPL 216 253	D. Decamp <i>et al.</i>	(ALEPH Collab.)					
IMAZATO	92	PRL 69 877	J. Imazato <i>et al.</i>	(KEK, INUS, TOKY+)					
MISHRA	92	PRL 68 3499	S.R. Mishra <i>et al.</i>	(COLU, CHIC, FNAL+)					
POLAK	92B	PR D46 3871	J. Polak, M. Zralek	(SILES)					
ACTON	91	PL B268 122	D.P. Acton <i>et al.</i>	(OPAL Collab.)					
ACTON	91B	PL B273 338	D.P. Acton <i>et al.</i>	(OPAL Collab.)					
ADEVA	91D	PL B262 159	B. Adeva <i>et al.</i>	(L3 Collab.)					
AQUINO	91	PL B261 280	M. Aquino, A. Fernandez, A. Garcia	(CINV, PUEB)					
COLANGELO	91	PL B253 154	P. Colangelo, G. Nardulli	(BARI)					
CUYPERS	91	PL B259 173	F. Cuyper, A.F. Falk, P.H. Frampton	(DURH, HARV+)					
FARAGGI	91	MPL A6 611	A.E. Faraggi, D.V. Nanopoulos	(TAMU)					
POLAK	91	NP B363 385	J. Polak, M. Zralek	(SILES)					

## Axions ( $A^0$ ) and Other Very Light Bosons, Searches for

### AXIONS AND OTHER SIMILAR PARTICLES

Revised January 2016 by A. Ringwald (DESY), L.J. Rosenberg and G. Rybka (U. of Washington).

#### Introduction

In this section, we list coupling-strength and mass limits for light neutral scalar or pseudoscalar bosons that couple weakly to normal matter and radiation. Such bosons may arise from a global spontaneously broken  $U(1)$  symmetry, resulting in a massless Nambu-Goldstone (NG) boson. If there is a small explicit symmetry breaking, either already in the Lagrangian or due to quantum effects such as anomalies, the boson acquires a mass and is called a pseudo-NG boson. Typical examples are axions ( $A^0$ ) [1,2], familons [3] and Majorons [4], associated, respectively, with a spontaneously broken Peccei-Quinn, family and lepton-number symmetry.

A common characteristic among these light bosons  $\phi$  is that their coupling to Standard-Model particles is suppressed by the energy scale that characterizes the symmetry breaking, *i.e.*, the decay constant  $f$ . The interaction Lagrangian is

$$\mathcal{L} = f^{-1} J^\mu \partial_\mu \phi, \quad (1)$$

where  $J^\mu$  is the Noether current of the spontaneously broken global symmetry. If  $f$  is very large, these new particles interact very weakly. Detecting them would provide a window to physics far beyond what can be probed at accelerators.

Axions are of particular interest because the Peccei-Quinn (PQ) mechanism remains perhaps the most credible scheme to preserve CP in QCD. Moreover, the cold dark matter of the universe may well consist of axions and they are searched for in dedicated experiments with a realistic chance of discovery.

Originally it was assumed that the PQ scale  $f_A$  was related to the electroweak symmetry-breaking scale  $v_{\text{weak}} =$

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$(\sqrt{2}G_F)^{-1/2} = 247$  GeV. However, the associated “standard” and “variant” axions were quickly excluded—we refer to the Listings for detailed limits. Here we focus on “invisible axions” with  $f_A \gg v_{\text{weak}}$  as the main possibility.

Axions have a characteristic two-photon vertex, inherited from their mixing with  $\pi^0$  and  $\eta$ . This coupling allows for the main search strategy based on axion-photon conversion in external magnetic fields [5], an effect that also can be of astrophysical interest. While for axions the product “ $A\gamma\gamma$  interaction strength  $\times$  mass” is essentially fixed by the corresponding  $\pi^0$  properties, one may consider a more general class of axion-like particles (ALPs) where the two parameters (coupling and mass) are independent. A number of experiments explore this more general parameter space. ALPs populating the latter are predicted to arise generically, in addition to the axion, in low-energy effective field theories emerging from string theory [6]. The latter often contain also very light Abelian vector bosons under which the Standard-Model particles are not charged: so-called hidden-sector photons, dark photons or paraphotons. They share a number of phenomenological features with the axion and ALPs, notably the possibility of hidden photon to photon conversion. Their physics cases and the current constraints are compiled in Ref. [7].

### I. THEORY

#### I.1 Peccei-Quinn mechanism and axions

The QCD Lagrangian includes a CP-violating term  $\mathcal{L}_\Theta = -\bar{\Theta}(\alpha_s/8\pi)G^{\mu\nu a}\tilde{G}_{\mu\nu}^a$ , where  $-\pi \leq \bar{\Theta} \leq +\pi$  is the effective  $\bar{\Theta}$  parameter after diagonalizing quark masses,  $G_{\mu\nu}^a$  is the color field strength tensor, and  $\tilde{G}^{a,\mu\nu} \equiv \epsilon^{\mu\nu\lambda\rho}G_{\lambda\rho}^a/2$ , with  $\epsilon^{0123} = 1$ , its dual. Limits on the neutron electric dipole moment [8] imply  $|\bar{\Theta}| \lesssim 10^{-10}$  even though  $\bar{\Theta} = \mathcal{O}(1)$  is otherwise completely satisfactory. The spontaneously broken global Peccei-Quinn symmetry  $U(1)_{\text{PQ}}$  was introduced to solve this “strong CP problem” [1], the axion being the pseudo-NG boson of  $U(1)_{\text{PQ}}$  [2]. This symmetry is broken due to the axion’s anomalous triangle coupling to gluons,

$$\mathcal{L} = \left( \frac{\phi_A}{f_A} - \bar{\Theta} \right) \frac{\alpha_s}{8\pi} G^{\mu\nu a} \tilde{G}_{\mu\nu}^a, \quad (2)$$

where  $\phi_A$  is the axion field and  $f_A$  the axion decay constant. Color anomaly factors have been absorbed in the normalization of  $f_A$  which is defined by this Lagrangian. Thus normalized,  $f_A$  is the quantity that enters all low-energy phenomena [9]. Non-perturbative topological fluctuations of the gluon fields in QCD induce a potential for  $\phi_A$  whose minimum is at  $\phi_A = \bar{\Theta} f_A$ , thereby canceling the  $\bar{\Theta}$  term in the QCD Lagrangian and thus restoring CP symmetry.

The resulting axion mass, in units of the PQ scale  $f_A$ , is identical to the square root of the topological susceptibility in QCD,  $m_A f_A = \sqrt{\chi}$ . The latter can be evaluated further [10], exploiting the chiral limit (masses of up and down quarks much smaller than the scale of QCD), yielding  $m_A f_A = \sqrt{\chi} \approx f_\pi m_\pi$ ,

where  $m_\pi = 135$  MeV and  $f_\pi \approx 92$  MeV. In more detail one finds, to leading order in chiral perturbation theory,

$$m_A = \frac{z^{1/2}}{1+z} \frac{f_\pi m_\pi}{f_A} = \frac{0.60 \text{ MeV}}{f_A/10^{10} \text{ GeV}}, \quad (3)$$

where  $z = m_u/m_d$ . We have used the canonical value  $z = 0.56$  [11], although the range  $z = 0.38\text{--}0.58$  is plausible [12]. The next-to-leading order correction to the axion mass has been evaluated recently in Ref. [13].

Originally one assumed  $f_A \sim v_{\text{weak}}$  [1,2]. Tree-level flavor conservation fixes the axion properties in terms of a single parameter: the ratio of the vacuum expectation values of two Higgs fields that appear as a minimal ingredient. This “standard axion” was excluded after extensive searches [14]. A narrow peak structure observed in positron spectra from heavy ion collisions [15] suggested an axion-like particle of mass 1.8 MeV that decays into  $e^+e^-$ , but extensive follow-up searches were negative. “Variant axion models” were proposed which keep  $f_A \sim v_{\text{weak}}$  while relaxing the constraint of tree-level flavor conservation [16], but these models are also excluded [17].

However, axions with  $f_A \gg v_{\text{weak}}$  evade all current experimental limits. One generic class of models invokes “hadronic axions” where new heavy quarks carry  $U(1)_{\text{PQ}}$  charges, leaving ordinary quarks and leptons without tree-level axion couplings. The archetype is the KSVZ model [18], where in addition the heavy quarks are electrically neutral. Another generic class requires at least two Higgs doublets and ordinary quarks and leptons carry PQ charges, the archetype being the DFSZ model [19]. All of these models contain at least one electroweak singlet scalar that acquires a vacuum expectation value and thereby breaks the PQ symmetry. The KSVZ and DFSZ models are frequently used as benchmark examples, but other models exist where both heavy quarks and Higgs doublets carry PQ charges. In supersymmetric models, the axion is part of a supermultiplet and thus inevitably accompanied by a spin-0 saxion and a spin-1 axino, which both also have couplings suppressed by  $f_A$ , and are expected to have large masses due to supersymmetry breaking [20].

#### I.2 Model-dependent axion couplings

Although the generic axion interactions scale approximately with  $f_\pi/f_A$  from the corresponding  $\pi^0$  couplings, there are non-negligible model-dependent factors and uncertainties. The axion’s two-photon interaction plays a key role for many searches,

$$\mathcal{L}_{A\gamma\gamma} = -\frac{G_{A\gamma\gamma}}{4} F_{\mu\nu} \tilde{F}^{\mu\nu} \phi_A = G_{A\gamma\gamma} \mathbf{E} \cdot \mathbf{B} \phi_A, \quad (4)$$

where  $F$  is the electromagnetic field-strength tensor and  $\tilde{F}^{\mu\nu} \equiv \epsilon^{\mu\nu\lambda\rho}F_{\lambda\rho}/2$ , with  $\epsilon^{0123} = 1$ , its dual. The coupling constant is

$$\begin{aligned} G_{A\gamma\gamma} &= \frac{\alpha}{2\pi f_A} \left( \frac{E}{N} - \frac{2}{3} \frac{4+z}{1+z} \right) \\ &= \frac{\alpha}{2\pi} \left( \frac{E}{N} - \frac{2}{3} \frac{4+z}{1+z} \right) \frac{1+z}{z^{1/2}} \frac{m_A}{m_\pi f_\pi}, \end{aligned} \quad (5)$$

where  $E$  and  $N$  are the electromagnetic and color anomalies of the axial current associated with the axion. In grand unified

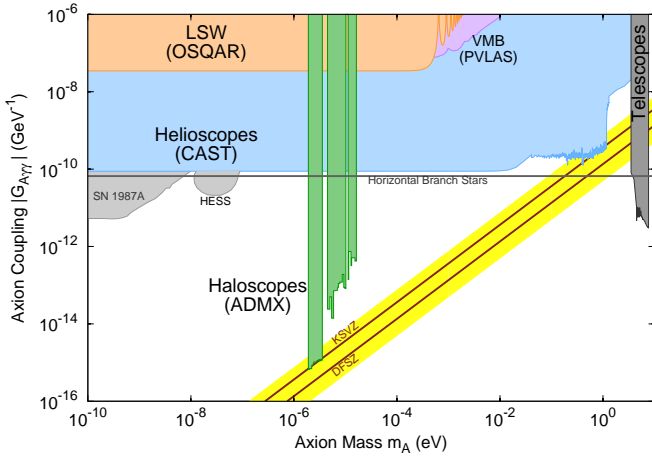
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models, and notably for DFSZ [19],  $E/N = 8/3$ , whereas for KSVZ [18]  $E/N = 0$  if the electric charge of the new heavy quark is taken to vanish. In general, a broad range of  $E/N$  values is possible [21], as indicated by the yellow band in Figure 1. The two-photon decay width is

$$\Gamma_{A \rightarrow \gamma\gamma} = \frac{G_{A\gamma\gamma}^2 m_A^3}{64\pi} = 1.1 \times 10^{-24} \text{ s}^{-1} \left( \frac{m_A}{\text{eV}} \right)^5. \quad (6)$$

The second expression uses Eq. (5) with  $z = 0.56$  and  $E/N = 0$ . Axions decay faster than the age of the universe if  $m_A \gtrsim 20 \text{ eV}$ .



**Figure 1:** Exclusion plot for axion-like particles as described in the text.

The interaction with fermions  $f$  has derivative form and is invariant under a shift  $\phi_A \rightarrow \phi_A + \phi_0$  as behooves a NG boson,

$$\mathcal{L}_{Aff} = \frac{C_f}{2f_A} \bar{\Psi}_f \gamma^\mu \gamma_5 \Psi_f \partial_\mu \phi_A. \quad (7)$$

Here,  $\Psi_f$  is the fermion field,  $m_f$  its mass, and  $C_f$  a model-dependent coefficient. The dimensionless combination  $g_{Aff} \equiv C_f m_f / f_A$  plays the role of a Yukawa coupling and  $\alpha_{Aff} \equiv g_{Aff}^2 / 4\pi$  of a “fine-structure constant.” The often-used pseudoscalar form  $\mathcal{L}_{Aff} = -i(C_f m_f / f_A) \bar{\Psi}_f \gamma_5 \Psi_f \phi_A$  need not be equivalent to the appropriate derivative structure, for example when two NG bosons are attached to one fermion line as in axion emission by nucleon bremsstrahlung [22].

In the DFSZ model [19], the tree-level coupling coefficient to electrons is [23]

$$C_e = \frac{\cos^2 \beta'}{3}, \quad (8)$$

where  $\tan \beta' = v_d / v_u$  is the ratio of the vacuum expectation value  $v_d$  of the Higgs field  $H_d$  giving masses to the down-type quarks and the vacuum expectation value  $v_u$  of the Higgs field  $H_u$  giving masses to the up-type quarks. (The prime at the angle indicates that the convention in the axion literature differs from the one in the Higgs literature, which uses  $\tan \beta = v_u / v_d = \cot \beta'$  [24].)

For nucleons,  $C_{n,p}$  are related to axial-vector current matrix elements by generalized Goldberger-Treiman relations,

$$\begin{aligned} C_p &= (C_u - \eta) \Delta u + (C_d - \eta z) \Delta d + (C_s - \eta w) \Delta s, \\ C_n &= (C_u - \eta) \Delta d + (C_d - \eta z) \Delta u + (C_s - \eta w) \Delta s. \end{aligned} \quad (9)$$

Here,  $\eta = (1 + z + w)^{-1}$  with  $z = m_u / m_d$  and  $w = m_u / m_s \ll z$  and the  $\Delta q$  are given by the axial vector current matrix element  $\Delta q S_\mu = \langle p | \bar{q} \gamma_\mu \gamma_5 q | p \rangle$  with  $S_\mu$  the proton spin.

Neutron beta decay and strong isospin symmetry considerations imply  $\Delta u - \Delta d = F + D = 1.269 \pm 0.003$ , whereas hyperon decays and flavor SU(3) symmetry imply  $\Delta u + \Delta d - 2\Delta s = 3F - D = 0.586 \pm 0.031$  [25]. The strange-quark contribution is  $\Delta s = -0.08 \pm 0.01_{\text{stat}} \pm 0.05_{\text{sys}}$  from the COMPASS experiment [26], and  $\Delta s = -0.085 \pm 0.008_{\text{exp}} \pm 0.013_{\text{theor}} \pm 0.009_{\text{evol}}$  from HERMES [25], in agreement with each other and with an early estimate of  $\Delta s = -0.11 \pm 0.03$  [27]. We thus adopt  $\Delta u = 0.84 \pm 0.02$ ,  $\Delta d = -0.43 \pm 0.02$  and  $\Delta s = -0.09 \pm 0.02$ , very similar to what was used in the axion literature.

The uncertainty of the axion-nucleon couplings is dominated by the uncertainty  $z = m_u / m_d = 0.38 - 0.58$  that we mentioned earlier. For hadronic axions  $C_{u,d,s} = 0$ , so that  $-0.51 < C_p < -0.36$  and  $0.10 > C_n > -0.05$ . Therefore it is well possible that  $C_n = 0$  whereas  $C_p$  does not vanish within the plausible  $z$  range. In the DFSZ model,  $C_u = \frac{1}{3} \sin^2 \beta'$  and  $C_d = \frac{1}{3} \cos^2 \beta'$  and  $C_n$  and  $C_p$  as functions of  $\beta'$  and  $z$  do not vanish simultaneously.

The axion-pion interaction is given by the Lagrangian [28]

$$\mathcal{L}_{A\pi} = \frac{C_{A\pi}}{f_\pi f_A} (\pi^0 \pi^+ \partial_\mu \pi^- + \pi^0 \pi^- \partial_\mu \pi^+ - 2\pi^+ \pi^- \partial_\mu \pi^0) \partial_\mu \phi_A, \quad (10)$$

where  $C_{A\pi} = (1 - z) / [3(1 + z)]$  in hadronic models. The chiral symmetry-breaking Lagrangian provides an additional term  $\mathcal{L}'_{A\pi} \propto (m_\pi^2 / f_\pi f_A) (\pi^0 \pi^0 + 2\pi^- \pi^+) \pi^0 \phi_A$ . For hadronic axions it vanishes identically, in contrast to the DFSZ model (Roberto Peccei, private communication).

## II. LABORATORY SEARCHES

### II.1 Light shining through walls

Searching for “invisible axions” is extremely challenging due to its extraordinarily feeble coupling to normal matter and radiation. Currently, the most promising approaches rely on the axion-two-photon vertex, allowing for axion-photon conversion in external electric or magnetic fields [5]. For the Coulomb field of a charged particle, the conversion is best viewed as a scattering process,  $\gamma + Ze \leftrightarrow Ze + A$ , called Primakoff effect [29]. In the other extreme of a macroscopic field, usually a large-scale  $B$ -field, the momentum transfer is small, the interaction coherent over a large distance, and the conversion is best viewed as an axion-photon oscillation phenomenon in analogy to neutrino flavor oscillations [30].

Photons propagating through a transverse magnetic field, with incident  $\mathbf{E}_\gamma$  and magnet  $\mathbf{B}$  parallel, may convert into axions. For  $m_A^2 L / 2\omega \ll 2\pi$ , where  $L$  is the length of the  $B$  field region and  $\omega$  the photon energy, the resultant axion beam is coherent with the incident photon beam and the



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conversion probability is  $\Pi \sim (1/4)(G_{A\gamma\gamma}BL)^2$ . A practical realization uses a laser beam propagating down the bore of a superconducting dipole magnet (like the bending magnets in high-energy accelerators). If another magnet is in line with the first, but shielded by an optical barrier, then photons may be regenerated from the pure axion beam [31]. The overall probability is  $P(\gamma \rightarrow A \rightarrow \gamma) = \Pi^2$ .

The first such experiment utilized two magnets of length  $L = 4.4$  m and  $B = 3.7$  T and found  $|G_{A\gamma\gamma}| < 6.7 \times 10^{-7} \text{ GeV}^{-1}$  at 95% CL for  $m_A < 1$  meV [32]. More recently, several such experiments were performed (see Listings) [33,34]. The current best limit,  $|G_{A\gamma\gamma}| < 3.5 \times 10^{-8} \text{ GeV}^{-1}$  at 95% CL for  $m_A \lesssim 0.3$  meV (see Figure 1), has been achieved by the OSQAR (Optical Search for QED Vacuum Birefringence, Axions, and Photon Regeneration) experiment, which exploited two 9 T LHC dipole magnets and an 18.5 W continuous wave laser emitting at the wavelength of 532 nm [34]. Some of these experiments have also reported limits for scalar bosons where the photon  $\mathbf{E}_\gamma$  must be chosen perpendicular to the magnet  $\mathbf{B}$ .

The concept of resonantly enhanced photon regeneration may open unexplored regions of coupling strength [35]. In this scheme, both the production and detection magnets are within Fabry-Perot optical cavities and actively locked in frequency. The  $\gamma \rightarrow A \rightarrow \gamma$  rate is enhanced by a factor  $2\mathcal{F}\mathcal{F}'/\pi^2$  relative to a single-pass experiment, where  $\mathcal{F}$  and  $\mathcal{F}'$  are the finesses of the two cavities. The resonant enhancement could be of order  $10^{(10-12)}$ , improving the  $G_{A\gamma\gamma}$  sensitivity by  $10^{(2.5-3)}$ . The experiment ALPS II (Any Light Particle Search II) is based on this concept and aims at an improvement of the current laboratory bound on  $G_{A\gamma\gamma}$  by a factor  $\sim 3 \times 10^3$  in the year 2018 [36].

Resonantly enhanced photon regeneration has already been exploited in experiments searching for "radiowaves shining through a shielding" [37,38]. For  $m_A \lesssim 10^{-5}$  eV, the upper bound on  $G_{A\gamma\gamma}$  established by the CROWS (CERN Resonant Weakly Interacting sub-eV Particle Search) experiment [39] is slightly less stringent than the one set by OSQAR.

### II.2 Photon polarization

An alternative to regenerating the lost photons is to use the beam itself to detect conversion: the polarization of light propagating through a transverse  $B$  field suffers dichroism and birefringence [40]. Dichroism: The  $E_\parallel$  component, but not  $E_\perp$ , is depleted by axion production, causing a small rotation of linearly polarized light. For  $m_A^2 L/2\omega \ll 2\pi$ , the effect is independent of  $m_A$ . For heavier axions, it oscillates and diminishes as  $m_A$  increases, and it vanishes for  $m_A > \omega$ . Birefringence: This rotation occurs because there is mixing of virtual axions in the  $E_\parallel$  state, but not for  $E_\perp$ . Hence, linearly polarized light will develop elliptical polarization. Higher-order QED also induces vacuum magnetic birefringence (VMB). A search for these effects was performed in the same dipole magnets in the early experiment above [41]. The dichroic rotation gave a stronger limit than the ellipticity rotation:

$|G_{A\gamma\gamma}| < 3.6 \times 10^{-7} \text{ GeV}^{-1}$  at 95% CL for  $m_A < 5 \times 10^{-4}$  eV. The ellipticity limits are better at higher masses, as they fall off smoothly and do not terminate at  $m_A$ .

In 2006 the PVLAS collaboration reported a signature of magnetically induced vacuum dichroism that could be interpreted as the effect of a pseudoscalar with  $m_A = 1\text{--}1.5$  meV and  $|G_{A\gamma\gamma}| = (1.6\text{--}5) \times 10^{-6} \text{ GeV}^{-1}$  [42]. Since then, these findings are attributed to instrumental artifacts [43]. This particle interpretation is also excluded by the above photon regeneration searches that were inspired by the original PVLAS result. Recently, the fourth generation setup of the PVLAS experiment has published new results on searches for VMB (see Figure 1) and dichroism [44]. The bounds from the non-observation of the latter on  $G_{A\gamma\gamma}$  are slightly weaker than the ones from OSQAR.

### II.3 Long-range forces

New bosons would mediate long-range forces, which are severely constrained by "fifth force" experiments [45]. Those looking for new mass-spin couplings provide significant constraints on pseudoscalar bosons [46]. Presently, the most restrictive limits are obtained from combining long-range force measurements with stellar cooling arguments [47]. For the moment, any of these limits are far from realistic values expected for axions. Still, these efforts provide constraints on more general low-mass bosons.

Recently, a method was proposed that can extend the search for axion-mediated spin-dependent forces by several orders of magnitude [48]. By combining techniques used in nuclear magnetic resonance and short-distance tests of gravity, this method appears to be sensitive to axions in the  $\mu\text{eV} - \text{meV}$  mass range, independent of the cosmic axion abundance.

## III. AXIONS FROM ASTROPHYSICAL SOURCES

### III.1 Stellar energy-loss limits:

Low-mass weakly-interacting particles (neutrinos, gravitons, axions, baryonic or leptonic gauge bosons, *etc.*) are produced in hot astrophysical plasmas, and can thus transport energy out of stars. The coupling strength of these particles with normal matter and radiation is bounded by the constraint that stellar lifetimes or energy-loss rates not conflict with observation [49–51].

We begin this discussion with our Sun and concentrate on hadronic axions. They are produced predominantly by the Primakoff process  $\gamma + Ze \rightarrow Ze + A$ . Integrating over a standard solar model yields the axion luminosity [52]

$$L_A = G_{10}^2 1.85 \times 10^{-3} L_\odot, \quad (11)$$

where  $G_{10} = |G_{A\gamma\gamma}| \times 10^{10} \text{ GeV}$ . The maximum of the spectrum is at 3.0 keV, the average at 4.2 keV, and the number flux at Earth is  $G_{10}^2 3.75 \times 10^{11} \text{ cm}^{-2} \text{ s}^{-1}$ . The solar photon luminosity is fixed, so axion losses require enhanced nuclear energy production and thus enhanced neutrino fluxes. The all-flavor measurements by SNO together with a standard solar

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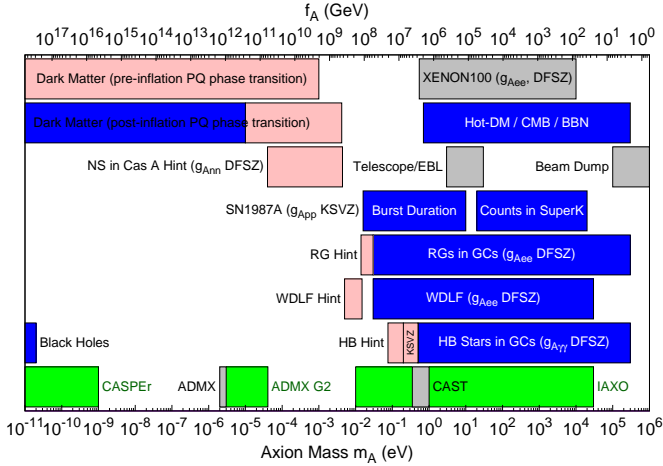
### Axions ( $A^0$ ) and Other Very Light Bosons

model imply  $L_A \lesssim 0.10 L_\odot$ , corresponding to  $G_{10} \lesssim 7$  [53], mildly superseding a similar limit from helioseismology [54]. Recently, the limit was improved to  $G_{10} < 4.1$  (at  $3\sigma$ ), exploiting a new statistical analysis that combined helioseismology (sound speed, surface helium and convective radius) and solar neutrino observations, including theoretical and observational errors, and accounting for tensions between input parameters of solar models, in particular the solar element abundances [55].

A more restrictive limit derives from globular-cluster (GC) stars that allow for detailed tests of stellar-evolution theory. The stars on the horizontal branch (HB) in the color-magnitude diagram have reached helium burning with a core-averaged energy release of about  $80 \text{ erg g}^{-1} \text{ s}^{-1}$ , compared to Primakoff axion losses of  $G_{10}^2 30 \text{ erg g}^{-1} \text{ s}^{-1}$ . The accelerated consumption of helium reduces the HB lifetime by about  $80/(80+30 G_{10}^2)$ . Number counts of HB stars in a large sample of 39 Galactic GCs compared with the number of red giants (that are not much affected by Primakoff losses) give a weak indication of non-standard losses which may be accounted by Primakoff-like axion emission, if the photon coupling is in the range  $|G_{A\gamma\gamma}| = 4.5^{+1.2}_{-1.6} \times 10^{-11} \text{ GeV}^{-1}$  [56]. Still, the upper bound found in this analysis,

$$|G_{A\gamma\gamma}| < 6.6 \times 10^{-11} \text{ GeV}^{-1} \text{ (95\% CL)}, \quad (12)$$

represents the strongest limit on  $G_{A\gamma\gamma}$  for a wide mass range, see Figure 1.



**Figure 2:** Exclusion ranges as described in the text. The intervals in the bottom row are the approximate ADMX, CASPER, CAST, and IAXO search ranges, with green regions indicating the projected reach. Limits on coupling strengths are translated into limits on  $m_A$  and  $f_A$  using  $z = 0.56$  and the KSVZ values for the coupling strengths, if not indicated otherwise. The “Beam Dump” bar is a rough representation of the exclusion range for standard or variant axions. The limits for the axion-electron coupling are determined for the DFSZ model with an axion-electron coupling corresponding to  $\cos^2 \beta' = 1/2$ .

We translate the conservative constraint, Equation 12, on  $G_{A\gamma\gamma}$  to  $f_A > 3.4 \times 10^7 \text{ GeV}$  ( $m_A < 0.2 \text{ eV}$ ), using  $z = 0.56$  and  $E/N = 0$  as in the KSVZ model, and show the excluded range in Figure 2. For the DFSZ model with  $E/N = 8/3$ , the corresponding limits are slightly less restrictive,  $f_A > 1.3 \times 10^7 \text{ GeV}$  ( $m_A < 0.5 \text{ eV}$ ). The weak indication of an extra energy loss points to a range  $76 \text{ meV} \lesssim m_A \lesssim 150 \text{ meV}$  ( $0.21 \text{ eV} \lesssim m_A \lesssim 0.41 \text{ eV}$ ) for the KSVZ (DFSZ) model. The exact high-mass end of the exclusion range has not been determined. The relevant temperature is around  $10 \text{ keV}$  and the average photon energy is therefore around  $30 \text{ keV}$ . The excluded  $m_A$  range thus certainly extends beyond the shown  $100 \text{ keV}$ .

If axions couple directly to electrons, the dominant emission processes are atomic axio-recombination and axio-deexcitation, axio-bremsstrahlung in electron-ion or electron-electron collisions, and Compton scattering [57]. Stars in the red giant (RG) branch of the color-magnitude diagram of GCs are particularly sensitive to these processes. In fact, they would lead to an extension of the latter to larger brightness. A recent analysis provided high-precision photometry for the Galactic globular cluster M5 (NGC 5904), allowing for a detailed comparison between the observed tip of the RG branch with predictions based on state-of-the-art stellar evolution theory [58]. It was found that, within the uncertainties, the observed and predicted tip of the RG branch brightness agree reasonably well within uncertainties, leading to the bound

$$\alpha_{Aee} < 1.5 \times 10^{-26} \text{ (95\% CL)}, \quad (13)$$

implying an upper bound on the axion mass in the DFSZ model,

$$m_A \cos^2 \beta' < 15 \text{ meV (95\% CL)}, \quad (14)$$

see Figure 2. Intriguingly, the agreement would improve with a small amount of extra cooling that slightly postpones helium ignition, preferring an electron coupling around  $\alpha_{Aee} \sim 2.8 \times 10^{-27}$ , corresponding to  $m_A \cos^2 \beta' \sim 7 \text{ meV}$ . Recently, it has been pointed out that the best fit simultaneously explaining the extra energy losses of HB stars reported above and the ones of RGs prefers a photon coupling around  $G_{A\gamma\gamma} \sim \text{few} \times 10^{-12} \text{ GeV}^{-1}$  and an electron coupling of order  $\alpha_{Aee} \sim 10^{-27}$  [59].

Bremsstrahlung is also efficient in white dwarfs (WDs), where the Primakoff and Compton processes are suppressed by the large plasma frequency. A comparison of the predicted and observed luminosity function of WDs can be used to put limits on  $\alpha_{Aee}$  [60]. A recent analysis, based on detailed WD cooling treatment and new data on the WD luminosity function (WDLF) of the Galactic Disk, found that electron couplings above  $\alpha_{Aee} \gtrsim 6 \times 10^{-27}$ , corresponding to a DFSZ axion mass  $m_A \cos^2 \beta' \gtrsim 10 \text{ meV}$ , are disfavoured [61], see Figure 2. Lower couplings can not be discarded from the current knowledge of the WDLF of the Galactic Disk. On the contrary, features in some WDLFs can be interpreted as suggestions for electron couplings in the range  $4.1 \times 10^{-28} \lesssim$

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$\alpha_{Aee} \lesssim 3.7 \times 10^{-27}$ , corresponding to  $2.5 \text{ MeV} \lesssim m_A \cos^2 \beta' \lesssim 7.5 \text{ MeV}$  [61,62], see Figure 2. For pulsationally unstable WDs (ZZ Ceti stars), the period decrease  $\dot{P}/P$  is a measure of the cooling speed. The corresponding observations of the pulsating WDs G117-B15A and R548 imply additional cooling that can be interpreted also in terms of similar axion losses [63].

Similar constraints derive from the measured duration of the neutrino signal of the supernova SN 1987A. Numerical simulations for a variety of cases, including axions and Kaluza-Klein gravitons, reveal that the energy-loss rate of a nuclear medium at the density  $3 \times 10^{14} \text{ g cm}^{-3}$  and temperature 30 MeV should not exceed about  $1 \times 10^{19} \text{ erg g}^{-1} \text{ s}^{-1}$  [50]. The energy-loss rate from nucleon bremsstrahlung,  $N + N \rightarrow N + N + A$ , is  $(C_N/2f_A)^2 (T^4/\pi^2 m_N) F$ . Here  $F$  is a numerical factor that represents an integral over the dynamical spin-density structure function because axions couple to the nucleon spin. For realistic conditions, even after considerable effort, one is limited to a heuristic estimate leading to  $F \approx 1$  [51].

The SN 1987A limits are of particular interest for hadronic axions where the bounds on  $\alpha_{Aee}$  are moot. Within uncertainties of  $z = m_u/m_d$  a reasonable choice for the coupling constants is then  $C_p = -0.4$  and  $C_n = 0$ . Using a proton fraction of 0.3,  $F = 1$ , and  $T = 30 \text{ MeV}$  one finds [51]

$$f_A \gtrsim 4 \times 10^8 \text{ GeV} \quad \text{and} \quad m_A \lesssim 16 \text{ MeV}, \quad (15)$$

see Figure 2. If axions interact sufficiently strongly they are trapped. Only about three orders of magnitude in  $g_{ANN}$  or  $m_A$  are excluded, a range shown somewhat schematically in Figure 2. For even larger couplings, the axion flux would have been negligible, yet it would have triggered additional events in the detectors, excluding a further range [64]. A possible gap between these two SN 1987A arguments was discussed as the “hadronic axion window” under the assumption that  $G_{A\gamma\gamma}$  was anomalously small [65]. This range is now excluded by hot dark matter bounds (see below).

Intriguingly, there is another hint for excessive stellar energy losses from the neutron star (NS) in the supernova remnant Cassiopeia A (Cas A): its surface temperature measured over 10 years reveals an unusually fast cooling rate. This may be interpreted as a hint for extra cooling by axion neutron bremsstrahlung, requiring a coupling to the neutron of size [66]

$$g_{Ann} = (3.8 \pm 3) \times 10^{-10} \quad (16)$$

corresponding to an axion mass

$$m_A = (2.4 \pm 2) \text{ MeV}/C_n, \quad (17)$$

see Figure 2. The hint is compatible with the state-of-the-art upper limit on this coupling,

$$g_{Ann} < 8 \times 10^{-10}, \quad (18)$$

from NS cooling [67]. In fact, as recently pointed out, the more rapid cooling of the superfluid core in the neutron star

may also arise from a phase transition of the neutron condensate into a multicomponent state [68].

Finally, let us note that if the interpretation of the various hints for additional cooling of stars reported in this section in terms of emission of axions with  $m_A \sim \text{MeV}$  were correct, SNe would lose a large fraction of their energy as axions. This would lead to a diffuse SN axion background in the universe with an energy density comparable to the extra-galactic background light [69]. However, there is no apparent way of detecting it or the axion burst from the next nearby SN.

### III.2 Searches for solar axions and ALPs

Instead of using stellar energy losses to derive axion limits, one can also search directly for these fluxes, notably from the Sun. The main focus has been on axion-like particles with a two-photon vertex. They are produced by the Primakoff process with a flux given by Equation 11 and an average energy of 4.2 keV, and can be detected at Earth with the reverse process in a macroscopic  $B$ -field (“axion helioscope”) [5]. In order to extend the sensitivity in mass towards larger values, one can endow the photon with an effective mass in a gas,  $m_\gamma = \omega_{\text{plas}}$ , thus matching the axion and photon dispersion relations [70].

An early implementation of these ideas used a conventional dipole magnet, with a conversion volume of variable-pressure gas with a xenon proportional chamber as x-ray detector [71]. The conversion magnet was fixed in orientation and collected data for about 1000 s/day. Axions were excluded for  $|G_{A\gamma\gamma}| < 3.6 \times 10^{-9} \text{ GeV}^{-1}$  for  $m_A < 0.03 \text{ eV}$ , and  $|G_{A\gamma\gamma}| < 7.7 \times 10^{-9} \text{ GeV}^{-1}$  for  $0.03 < m_A < 0.11 \text{ eV}$  at 95% CL.

Later, the Tokyo axion helioscope used a superconducting magnet on a tracking mount, viewing the Sun continuously. They reported  $|G_{A\gamma\gamma}| < 6 \times 10^{-10} \text{ GeV}^{-1}$  for  $m_A < 0.3 \text{ eV}$  [72]. This experiment was recommissioned and a similar limit for masses around 1 eV was reported [73].

The most recent helioscope CAST (CERN Axion Solar Telescope) uses a decommissioned LHC dipole magnet on a tracking mount. The hardware includes grazing-incidence x-ray optics with solid-state x-ray detectors, as well as a novel x-ray Micromegas position-sensitive gaseous detector. CAST has established a 95% CL limit  $|G_{A\gamma\gamma}| < 8.8 \times 10^{-11} \text{ GeV}^{-1}$  for  $m_A < 0.02 \text{ eV}$  [52]. To cover larger masses, the magnet bores are filled with a gas at varying pressure. The runs with  $^4\text{He}$  cover masses up to about 0.4 eV [74], providing the  $^4\text{He}$  limits shown in Figure 1. To cover yet larger masses,  $^3\text{He}$  was used to achieve a larger pressure at cryogenic temperatures. Limits up to 1.17 eV allowed CAST to “cross the axion line” for the KSVZ model [75], see Figure 1.

Recently, the XENON100 experiment has presented first results of searches for solar axions and ALPs [76]. The axion-electron coupling constant,  $g_{Aee}$ , has been probed by exploiting the axio-electric effect in liquid xenon, resulting in an upper bound

$$g_{Aee} < 7.7 \times 10^{-12} \quad (90\% \text{ CL}), \quad (19)$$

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excluding DFSZ models with  $m_A \cos^2 \beta' > 0.27 \text{ eV}$ , cf. see Figure 2.

Going to yet larger masses in a helioscope search is not well motivated because of the cosmic hot dark matter bound of  $m_A \lesssim 1 \text{ eV}$  (see below). Sensitivity to significantly smaller values of  $G_{A\gamma\gamma}$  can be achieved with a next-generation axion helioscope with a much larger magnetic-field cross section. Realistic design options for this “International Axion Observatory” (IAXO) have been studied in some detail [77]. Such a next-generation axion helioscope may also push the sensitivity in the product of couplings to photons and to electrons,  $G_{A\gamma\gamma} g_{Aee}$ , into a range beyond stellar energy-loss limits and test the hypothesis that WD cooling is dominated by axion emission [78].

Other Primakoff searches for solar axions and ALPs have been carried out using crystal detectors, exploiting the coherent conversion of axions into photons when the axion angle of incidence satisfies a Bragg condition with a crystal plane [79]. However, none of these limits is more restrictive than the one derived from the constraint on the solar axion luminosity ( $L_A \lesssim 0.10 L_\odot$ ) discussed earlier.

Another idea is to look at the Sun with an x-ray satellite when the Earth is in between. Solar axions and ALPs would convert in the Earth magnetic field on the far side and could be detected [80]. The sensitivity to  $G_{A\gamma\gamma}$  could be comparable to CAST, but only for much smaller  $m_A$ . Deep solar x-ray measurements with existing satellites, using the solar magnetosphere as conversion region, have reported preliminary limits on  $G_{A\gamma\gamma}$  [81].

#### III.3 Conversion of astrophysical photon fluxes

Large-scale  $B$  fields exist in astrophysics that can induce axion-photon oscillations. In practical cases,  $B$  is much smaller than in the laboratory, whereas the conversion region  $L$  is much larger. Therefore, while the product  $BL$  can be large, realistic sensitivities are usually restricted to very low-mass particles, far away from the “axion band” in a plot like Figure 1.

One example is SN 1987A, which would have emitted a burst of axion-like particles (ALPs) due to the Primakoff production in its core. They would have partially converted into  $\gamma$ -rays in the galactic  $B$ -field. The lack of a gamma-ray signal in the GRS instrument of the SMM satellite in coincidence with the observation of the neutrinos emitted from SN1987A therefore provides a strong bound on their coupling to photons [82]. Recently, this bound has been revisited and the underlying physics has been brought to the current state-of-the-art, as far as modelling of the supernova and the Milky-Way magnetic field are concerned, resulting in the limit [83]

$$|G_{A\gamma\gamma}| < 5.3 \times 10^{-12} \text{ GeV}^{-1}, \text{ for } m_A \lesssim 4.4 \times 10^{-10} \text{ eV}.$$

Magnetically induced oscillations between photons and axion-like particles (ALPs) can modify the photon fluxes from distant sources in various ways, featuring (i) frequency-dependent dimming, (ii) modified polarization, and (iii) avoid- ing absorption by propagation in the form of axions.

For example, dimming of SNe Ia could influence the interpretation in terms of cosmic acceleration [84], although it has become clear that photon-ALP conversion could only be a sub-dominant effect [85]. Searches for linearly polarised emission from magnetised white dwarfs [86] and changes of the linear polarisation from radio galaxies (see, e.g., Ref. [87]) provide limits close to  $G_{A\gamma\gamma} \sim 10^{-11} \text{ GeV}^{-1}$ , for masses  $m_A \lesssim 10^{-7} \text{ eV}$  and  $m_A \lesssim 10^{-15} \text{ eV}$ , respectively, albeit with uncertainties related to the underlying assumptions. Even stronger limits,  $G_{A\gamma\gamma} \lesssim 2 \times 10^{-13} \text{ GeV}^{-1}$ , for  $m_A \lesssim 10^{-14} \text{ eV}$ , have been obtained by exploiting high-precision measurements of quasar polarisations [88].

Remarkably, it appears that the universe could be too transparent to TeV  $\gamma$ -rays that should be absorbed by pair production on the extra-galactic background light [89]. The situation is not conclusive at present [90], but the possible role of photon-ALP oscillations in TeV  $\gamma$ -ray astronomy is tantalizing [91]. Fortunately, the region in ALP parameter space,  $G_{A\gamma\gamma} \sim 10^{-12} - 10^{-10} \text{ GeV}^{-1}$  for  $m_A \lesssim 10^{-7} \text{ eV}$  [92], required to explain the anomalous TeV transparency of the universe, could be conceivably probed by the next generation of laboratory experiments (ALPS II) and helioscopes (IAXO) mentioned above. This parameter region can also be probed by searching for an irregular behavior of the gamma ray spectrum of distant active galactic nuclei (AGN), expected to arise from photon-ALP mixing in a limited energy range. The H.E.S.S. collaboration has set a limit of  $|G_{A\gamma\gamma}| \lesssim 2.1 \times 10^{-11} \text{ GeV}^{-1}$ , for  $1.5 \times 10^{-8} \text{ eV} \lesssim m_A \lesssim 6.0 \times 10^{-8} \text{ eV}$ , from the non-observation of an irregular behavior of the spectrum of the AGN PKS 2155 [93], see Figure 1.

Last but not least, it was found that observed soft X-ray excesses in many galaxy clusters may be explained by the conversion of a hypothetical cosmic ALP background (CAB) radiation, corresponding to an effective number  $\Delta N_{\text{eff}}$  of extra neutrinos, into photons in the cluster magnetic fields [94]. This explanation requires that the CAB spectrum is peaked in the soft X-ray region and that the ALP coupling and mass satisfy  $|G_{A\gamma\gamma}| \gtrsim (1 - 2) \times 10^{-13} \text{ GeV}^{-1} \sqrt{0.5/\Delta N_{\text{eff}}}$ , for  $m_A \lesssim 10^{-12} \text{ eV}$ .

#### III.4 Superradiance of black holes

Ultralight bosonic fields such as axions or ALPs can affect the dynamics and gravitational wave emission of rapidly rotating astrophysical black holes through the Penrose superradiance mechanism. When their Compton wavelength is of order of the black hole size, they form bound states around the black hole nucleus. Their occupation number grows exponentially by extracting rotational energy and angular momentum from the ergosphere, thus forming a rotating Bose-Einstein condensate emitting gravitational waves. For black holes lighter than  $10^7$  solar masses, accretion cannot replenish the spin of the black hole. The existence of destabilizing ultralight bosonic fields thus leads to gaps in the mass vs. spin plot of rapidly rotating black holes. Stellar black hole spin measurements – exploiting

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well-studied binaries and two independent techniques – exclude a mass range  $6 \times 10^{-13} \text{ eV} < m_A < 2 \times 10^{-11} \text{ eV}$  at  $2\sigma$ , which for the axion excludes  $3 \times 10^{17} \text{ GeV} < f_A < 1 \times 10^{19} \text{ GeV}$  [95]. Long lasting, monochromatic gravitational wave signals, which can be distinguished from ordinary astrophysical sources, are expected to be produced by axions transitioning between the levels of the gravitational atom and axions annihilating to gravitons. Accordingly, the gravitational wave detector Advanced LIGO should be sensitive to the axion in the  $m_A \sim 10^{-10} \text{ eV}$  region.

#### IV. COSMIC AXIONS

##### IV.1 Cosmic axion populations

In the early universe, axions are produced by processes involving quarks and gluons [96]. After color confinement, the dominant thermalization process is  $\pi + \pi \leftrightarrow \pi + A$  [28]. The resulting axion population would contribute a hot dark matter component in analogy to massive neutrinos. Cosmological precision data provide restrictive constraints on a possible hot dark-matter fraction that translate into  $m_A \lesssim 1 \text{ eV}$  [97], but in detail depend on the used data set and assumed cosmological model. In the future, data from a EUCLID-like survey combined with Planck CMB data can detect hot dark matter axions mass  $m_A \gtrsim 0.15 \text{ eV}$  at very high significance [98].

For  $m_A \gtrsim 20 \text{ eV}$ , axions decay fast on a cosmic time scale, removing the axion population while injecting photons. This excess radiation provides additional limits up to very large axion masses [99]. An anomalously small  $G_{A\gamma\gamma}$  provides no loophole because suppressing decays leads to thermal axions overwhelming the mass density of the universe.

The main cosmological interest in axions derives from their possible role as cold dark matter (CDM). In addition to thermal processes, axions are abundantly produced by the “re-alignment mechanism” [100]. After the breakdown of the PQ symmetry, the axion field relaxes somewhere in the “bottom of the wine bottle” potential. Near the QCD epoch, topological fluctuations of the gluon fields such as instantons explicitly break the PQ symmetry, the very effect that causes dynamical PQ symmetry restoration. This “tilting of the wine bottle” drives the axion field toward the CP-conserving minimum, thereby exciting coherent oscillations of the axion field that ultimately represent a condensate of CDM. The fractional cosmic mass density in this homogeneous field mode is [101,102],

$$\begin{aligned} \Omega_A^{\text{real}} h^2 &\approx 0.11 \left( \frac{f_A}{5 \times 10^{11} \text{ GeV}} \right)^{1.19} F \bar{\Theta}_i^2 \\ &= 0.11 \left( \frac{12 \text{ } \mu\text{eV}}{m_A} \right)^{1.19} F \bar{\Theta}_i^2, \end{aligned} \quad (20)$$

where  $h$  is the present-day Hubble expansion parameter in units of  $100 \text{ km s}^{-1} \text{ Mpc}^{-1}$ , and  $-\pi \leq \bar{\Theta}_i \leq \pi$  is the initial “misalignment angle” relative to the CP-conserving position attained in the causally connected region which evolved into today’s observable universe.  $F = F(\bar{\Theta}_i, f_A)$  is a factor accounting for anharmonicities in the axion potential.

For  $F \bar{\Theta}_i^2 = \mathcal{O}(1)$ ,  $m_A$  should be above  $\sim 10 \text{ } \mu\text{eV}$  in order that the cosmic axion density does not exceed the observed CDM density,  $\Omega_{\text{CDM}} h^2 = 0.11$ . However, much smaller axion masses (much higher PQ scales) would still be possible if the PQ symmetry is broken during inflation and not restored afterwards. In this case, the initial value  $\bar{\Theta}_i$  may just happen to be small enough in today’s observable universe (“anthropic axion window” [103]). However, since the axion field is then present during inflation and thus subject to quantum fluctuations, the non-observation of the associated isocurvature fluctuations in the CMB puts severe constraints in the  $(f_A, r)$  plane, where  $r$  is the ratio of the power in tensor to the one in scalar fluctuations [104]. In fact, isocurvature constraints, combined with a future measurement of a sizeable  $r$ , would strongly disfavor axions with [105]

$$f_A \gtrsim 1.3 \times 10^{13} \text{ GeV} \left( \frac{r}{0.1} \right)^{1/2}, \quad m_A \lesssim 0.4 \text{ } \mu\text{eV} \left( \frac{r}{0.1} \right)^{-1/2}.$$

If the PQ symmetry breakdown takes place after inflation,  $\bar{\Theta}_i$  will take on different values in different patches of the universe. The average contribution is [101]

$$\Omega_A^{\text{real}} h^2 \approx 0.11 \left( \frac{41 \text{ } \mu\text{eV}}{m_A} \right)^{1.19}. \quad (21)$$

However, the additional contribution from the decay of topological defects suffers from significant uncertainties. According to Sikivie and collaborators, these populations are comparable to the re-alignment contribution [106]. Other groups find a significantly enhanced axion density [102,107] or rather, a larger  $m_A$  value for axions providing CDM, namely

$$m_A \approx (0.8 - 1.3) \times 10^{-4} \text{ eV}, \quad (22)$$

for models with short-lived (requiring unit color anomaly  $N = 1$ ) domain walls, such as the KSVZ model, and

$$m_A \approx (6 \times 10^{-4} - 4 \times 10^{-3}) \text{ eV}, \quad (23)$$

for models with long-lived ( $N > 1$ ) domain walls, such as an accidental DFSZ model [108], where the PQ symmetry is broken by higher dimensional Planck suppressed operators, see Figure 2. Moreover, the spatial axion density variations are large at the QCD transition and they are not erased by free streaming. When matter begins to dominate the universe, gravitationally bound “axion mini clusters” form promptly [109]. A significant fraction of CDM axions can reside in these bound objects.

In the above predictions of the fractional cosmic mass density in axions, the exponent, 1.19, arises from the non-trivial temperature dependence of the axion mass  $m_A(T) = \sqrt{\chi(T)}/f_A$ , which has been obtained from the dilute instanton gas/liquid approximation (DIGA). Lattice QCD provides a first principle technique to determine the topological susceptibility  $\chi(T)$  in the relevant temperature range around the QCD phase transition. A full result needs two ingredients: physical quark masses and a controlled continuum extrapolation from

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non-vanishing to zero lattice spacings. The latter has been done recently in the quenched framework (neglecting the effects of light quarks) and compared with the prediction of the DIGA [110,111]. Nice agreement was found for the temperature dependence, whereas the overall normalization of the DIGA result turned out to differ from the non-perturbative continuum extrapolated lattice results by a factor of order ten [111]. If this finding can be extrapolated to full QCD, the prediction of the axion mass relevant for dark matter will decrease by about 20% compared to the DIGA prediction. Lattice simulations with physical quark masses are about two-to-three orders of magnitude more CPU intensive than quenched ones. In addition one expects much smaller topological susceptibilities and larger cutoff effects. Correspondingly, available pioneering studies in full QCD [112] do not extend to the relevant temperature range and may still suffer from strong cutoff effects. But lattice campaigns dedicated to axion cosmology are ongoing.

In R-parity conserving supersymmetric models, more possibilities arise: cold dark matter might be a mixture of axions along with the lightest SUSY particle (LSP) [20]. Candidates for the LSP include the lightest neutralino, the gravitino, the axino, or a sneutrino. In the case of a neutralino LSP, saxion and axino production in the early universe have a strong impact on the neutralino and axion abundance. The former almost always gets increased beyond its thermal-production-only value, favoring then models with higgsino-like or wino-like neutralinos [113]. For large values of  $f_A$ , saxions from the vacuum re-alignment mechanism may produce large relic dilution via entropy dumping, thus allowing for much larger values of  $f_A$ , sometimes as high as approaching the GUT scale,  $\sim 10^{16}$  GeV, for natural values of the initial re-alignment angle. Then the dark matter may be either neutralino- or axion-dominated, or a comparable mixture. In such scenarios, one might expect eventual direct detection of both relic neutralinos and relic axions.

Finally, it is worth mentioning that the non-thermal production mechanisms attributed to axions are indeed generic to bosonic weakly interacting ultra-light particles such as ALPs: a wide range in  $G_{A\gamma\gamma} - m_A$  parameter space outside the axion band can generically contain models with adequate CDM density [114].

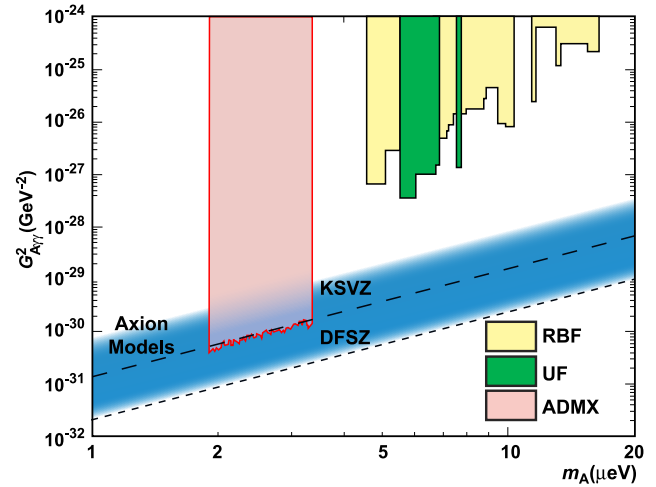
### IV.2 Telescope searches

The two-photon decay is extremely slow for axions with masses in the CDM regime, but could be detectable for eV masses. The signature would be a quasi-monochromatic emission line from galaxies and galaxy clusters. The expected optical line intensity for DFSZ axions is similar to the continuum night emission. An early search in three rich Abell clusters [115], and a recent search in two rich Abell clusters [116], exclude the “Telescope” range in Figure 1 and Figure 2 unless the axion-photon coupling is strongly suppressed. Of course, axions in this mass range would anyway provide an excessive hot DM contribution.

Very low-mass axions in halos produce a weak quasi-monochromatic radio line. Virial velocities in undisrupted dwarf galaxies are very low, and the axion decay line would therefore be extremely narrow. A search with the Haystack radio telescope on three nearby dwarf galaxies provided a limit  $|G_{A\gamma\gamma}| < 1.0 \times 10^{-9} \text{ GeV}^{-1}$  at 96% CL for  $298 < m_A < 363 \mu\text{eV}$  [117]. However, this combination of  $m_A$  and  $G_{A\gamma\gamma}$  does not exclude plausible axion models.

### IV.3 Microwave cavity experiments

The limits of Figure 2 suggest that axions, if they exist, provide a significant fraction or even perhaps all of the cosmic CDM. In a broad range of the plausible  $m_A$  range for CDM, galactic halo axions may be detected by their resonant conversion into a quasi-monochromatic microwave signal in a high-Q electromagnetic cavity permeated by a strong static  $B$  field [5,118]. The cavity frequency is tunable, and the signal is maximized when the frequency is the total axion energy, rest mass plus kinetic energy, of  $\nu = (m_A/2\pi) [1 + \mathcal{O}(10^{-6})]$ , the width above the rest mass representing the virial distribution in the galaxy. The frequency spectrum may also contain finer structure from axions more recently fallen into the galactic potential and not yet completely virialized [119].



**Figure 3:** Exclusion region reported from the microwave cavity experiments RBF and UF [120] and ADMX [121]. A local dark-matter density of  $450 \text{ MeV cm}^{-3}$  is assumed.

The feasibility of this technique was established in early experiments of relatively small sensitive volume,  $\mathcal{O}(1 \text{ liter})$ , with HFET-based amplifiers, setting limits in the range  $4.5 < m_A < 16.3 \mu\text{eV}$  [120], but lacking by 2–3 orders of magnitude the sensitivity required to detect realistic axions. Later, ADMX ( $B \sim 8 \text{ T}$ ,  $V \sim 200 \text{ liters}$ ) has achieved sensitivity to KSVZ axions, assuming they saturate the local dark matter density and are well virialized, over the mass range  $1.9\text{--}3.3 \mu\text{eV}$  [121]. Should halo axions have a significant component not yet virialized, ADMX is sensitive to DFSZ

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axions [122]. The corresponding 90% CL exclusion regions shown in Figure 3 are normalized to an assumed local CDM density of  $7.5 \times 10^{-25} \text{ g cm}^{-3}$  ( $450 \text{ MeV cm}^{-3}$ ). More recently the ADMX experiment commissioned an upgrade [123] that replaces the microwave HFET amplifiers by near quantum-limited low-noise dc SQUID microwave amplifiers [124], allowing for a significantly improved sensitivity [125]. This apparatus is also sensitive to other hypothetical light bosons, such as hidden photons or chameleons, over a limited parameter space [114,126]. Alternatively, a Rydberg atom single-photon detector [127] can in principle evade the standard quantum limit for coherent photon detection.

Other new concepts for searching for axion dark matter are also being investigated. For instance, photons from dark matter axions or ALPs could be focused in a manner similar to a dish antenna instead of a resonant cavity [128], enabling broadband searches at higher masses than the RF technique. Searches for hidden photon dark matter exploiting this technique are already underway [129]. Another alternative to the microwave cavity technique is based on a novel detector architecture consisting of an open, Fabry-Perot resonator and a series of current-carrying wire planes [130]. The Orpheus detector has demonstrated this new technique, excluding dark matter ALPs with masses between  $68.2$  and  $76.5 \mu\text{eV}$  and axion-photon couplings greater than  $4 \times 10^{-7} \text{ GeV}^{-1}$ . This technique may be able to probe dark matter axions in the mass range from  $40$  to  $700 \mu\text{eV}$ . Another proposed axion dark matter search method sensitive in the  $100 \mu\text{eV}$  mass range is to cool a kilogram-sized sample to millikelvin temperatures and count axion induced atomic transitions using laser techniques [131].

#### IV.4 Magnetic resonance searches

The oscillating galactic dark matter axion field induces oscillating nuclear electric dipole moments (EDMs). These EDMs cause the precession of nuclear spins in a nucleon spin polarized sample in the presence of an electric field. The resulting transverse magnetization can be searched for by exploiting magnetic-resonance (MR) techniques, which are most sensitive in the range of low oscillation frequencies corresponding to sub-neV axion masses. The aim of the corresponding Cosmic Axion Spin Precession Experiment (CASPER) [132] is to probe axion dark matter in the anthropic window,  $f_A \gtrsim 10^{15} \text{ GeV}$ , corresponding to  $m_A \lesssim \text{neV}$ , complementary to the classic axion window probed by the RF cavity technique.

In the intermediate mass region,  $\text{neV} \lesssim m_A \lesssim 0.1 \mu\text{eV}$ , one may exploit a cooled LC circuit and precision magnetometry to search for the oscillating electric current induced by dark matter axions in a strong magnetic field [133].

An eventually non-zero axion electron coupling  $g_{Aee}$  will lead to a spin precession about the axion dark matter wind [134]. The QUAX (QUaerere AXions) experiment aims at exploiting MR inside a magnetized material [135]. Because of the higher Larmor frequency of the electron, it is sensitive in the classic window.

#### Conclusions

There is a strengthening physics case for very weakly coupled ultralight particles beyond the Standard Model. The elegant solution of the strong CP problem proposed by Peccei and Quinn yields a particularly strong motivation for the axion. In many theoretically appealing ultraviolet completions of the Standard Model axions and axion-like particles occur automatically. Moreover, they are natural cold dark matter candidates. Perhaps the first hints of their existence have already been seen in the anomalous excessive cooling of stars and the anomalous transparency of the Universe for VHE gamma rays. Interestingly, a significant portion of previously unexplored, but phenomenologically very interesting and theoretically very well motivated axion and ALP parameter space can be tackled in the foreseeable future by a number of terrestrial experiments searching for axion/ALP dark matter, for solar axions/ALPs, and for light apparently shining through a wall.

#### References

1. R.D. Peccei and H. Quinn, Phys. Rev. Lett. **38**, 1440 (1977); Phys. Rev. **D16**, 1791 (1977).
2. S. Weinberg, Phys. Rev. Lett. **40**, 223 (1978); F. Wilczek, Phys. Rev. Lett. **40**, 279 (1978).
3. F. Wilczek, Phys. Rev. Lett. **49**, 1549 (1982).
4. Y. Chikashige, R.N. Mohapatra, and R.D. Peccei, Phys. Lett. **B98**, 265 (1981); G.B. Gelmini and M. Roncadelli, Phys. Lett. **B99**, 411 (1981).
5. P. Sikivie, Phys. Rev. Lett. **51**, 1415 (1983) and Erratum *ibid.*, **52**, 695 (1984).
6. E. Witten, Phys. Lett. **B149**, 351 (1984); J.P. Conlon, JHEP **0605**, 078 (2006); K.-S. Choi *et al.*, Phys. Lett. **B675**, 381 (2009); A. Arvanitaki *et al.*, Phys. Rev. **D81**, 123530 (2010); B.S. Acharya, K. Bobkov, and P. Kumar, JHEP **1011**, 105 (2010); M. Cicoli, M. Goodsell, and A. Ringwald, JHEP **1210**, 146 (2012).
7. J. Jaeckel and A. Ringwald, Ann. Rev. Nucl. and Part. Sci. **60**, 405 (2010); A. Ringwald, Phys. Dark Univ. **1**, 116 (2012); J. Jaeckel, Frascati Phys. Ser. **56**, 172 (2013).
8. C.A. Baker *et al.*, Phys. Rev. Lett. **97**, 131801 (2006).
9. H. Georgi, D.B. Kaplan, and L. Randall, Phys. Lett. **B169**, 73 (1986).
10. R.J. Crewther, Phys. Lett. **B70**, 349 (1977); P. Di Vecchia and G. Veneziano, Nucl. Phys. **B171**, 253 (1980).
11. H. Leutwyler, Phys. Lett. **B378**, 313 (1996).
12. Mini review on Quark Masses in: K.A. Olive *et al.* (Particle Data Group), Chin. Phys. C **38**, 090001 (2014).
13. G.G. di Cortona *et al.*, JHEP **1601**, 034 (2016).
14. T.W. Donnelly *et al.*, Phys. Rev. **D18**, 1607 (1978); S. Barshay *et al.*, Phys. Rev. Lett. **46**, 1361 (1981); A. Barroso and N.C. Mukhopadhyay, Phys. Lett. **B106**, 91 (1981); R.D. Peccei, in *Proceedings of Neutrino '81*, Honolulu, Hawaii, Vol. 1, p. 149 (1981);

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

- L.M. Krauss and F. Wilczek, Phys. Lett. **B173**, 189 (1986).
15. J. Schweppe *et al.*, Phys. Rev. Lett. **51**, 2261 (1983);  
T. Cowan *et al.*, Phys. Rev. Lett. **54**, 1761 (1985).
16. R.D. Peccei, T.T. Wu, and T. Yanagida, Phys. Lett. **B172**, 435 (1986).
17. W.A. Bardeen, R.D. Peccei, and T. Yanagida, Nucl. Phys. **B279**, 401 (1987).
18. J.E. Kim, Phys. Rev. Lett. **43**, 103 (1979);  
M.A. Shifman, A.I. Vainshtein, and V.I. Zakharov, Nucl. Phys. **B166**, 493 (1980).
19. M. Dine, W. Fischler, and M. Srednicki, Phys. Lett. **B104**, 199 (1981);  
A.R. Zhitnitsky, Sov. J. Nucl. Phys. **31**, 260 (1980).
20. J.E. Kim and G. Carosi, Rev. Mod. Phys. **82**, 557 (2010).
21. J.E. Kim, Phys. Rev. **D58**, 055006 (1998).
22. G. Raffelt and D. Seckel, Phys. Rev. Lett. **60**, 1793 (1988);  
M. Carena and R.D. Peccei, Phys. Rev. **D40**, 652 (1989);  
K. Choi, K. Kang, and J.E. Kim, Phys. Rev. Lett. **62**, 849 (1989).
23. M. Srednicki, Nucl. Phys. **B260**, 689 (1985).
24. Status of Higgs boson physics in: K.A. Olive *et al.* (Particle Data Group), Chin. Phys. C **38**, 090001 (2014).
25. A. Airapetian *et al.* (HERMES Collab.), Phys. Rev. **D75**, 012007 (2007) and Erratum *ibid.*, **D76**, 039901 (2007).
26. V.Y. Alexakhin *et al.* (COMPASS Collab.), Phys. Lett. **B647**, 8 (2007).
27. J.R. Ellis and M. Karliner, in: *The spin structure of the nucleon: International school of nucleon structure* (3–10 August 1995, Erice, Italy), ed. by B. Frois, V.W. Hughes, and N. De Groot (World Scientific, Singapore, 1997) [hep-ph/9601280].
28. S. Chang and K. Choi, Phys. Lett. **B316**, 51 (1993).
29. D.A. Dicus *et al.*, Phys. Rev. **D18**, 1829 (1978).
30. G. Raffelt and L. Stodolsky, Phys. Rev. **D37**, 1237 (1988).
31. A.A. Anselm, Yad. Fiz. **42**, 1480 (1985);  
K. van Bibber *et al.*, Phys. Rev. Lett. **59**, 759 (1987).
32. G. Ruoso *et al.*, Z. Phys. **C56**, 505 (1992);  
R. Cameron *et al.*, Phys. Rev. **D47**, 3707 (1993).
33. M. Fouche *et al.* (BMV Collab.), Phys. Rev. **D78**, 032013 (2008);  
P. Pugnat *et al.* (OSQAR Collab.), Phys. Rev. **D78**, 092003 (2008);  
A. Chou *et al.* (GammeV T-969 Collab), Phys. Rev. Lett. **100**, 080402 (2008);  
A. Afanasev *et al.* (LIPSS Collab.), Phys. Rev. Lett. **101**, 120401 (2008);  
K. Ehret *et al.* (ALPS Collab.), Phys. Lett. **B689**, 149 (2010);  
P. Pugnat *et al.* (OSQAR Collab.), Eur. Phys. J. **C74**, 3027 (2014).
34. R. Ballou *et al.* (OSQAR Collab.), Phys. Rev. **D92**, 092002 (2015).
35. F. Hoogeveen and T. Ziegenhagen, Nucl. Phys. **B358**, 3 (1991);  
P. Sikivie, D. Tanner, and K. van Bibber, Phys. Rev. Lett. **98**, 172002 (2007);  
G. Mueller *et al.*, Phys. Rev. **D80**, 072004 (2009).
36. R. Baehre *et al.* (ALPS Collab.), JINST **1308**, T09001 (2013).
37. F. Hoogeveen, Phys. Lett. **B288**, 195 (1992);  
J. Jaeckel and A. Ringwald, Phys. Lett. **B659**, 509 (2008);  
F. Caspers, J. Jaeckel, and A. Ringwald, JINST **0904**, P11013 (2009).
38. R. Povey, J. Hartnett, and M. Tobar, Phys. Rev. **D82**, 052003 (2010);  
A. Wagner *et al.*, Phys. Rev. Lett. **105**, 171801 (2010).
39. M. Betz *et al.*, Phys. Rev. **D88**, 075014 (2013).
40. L. Maiani *et al.*, Phys. Lett. **B175**, 359 (1986).
41. Y. Semertzidis *et al.*, Phys. Rev. Lett. **64**, 2988 (1990).
42. E. Zavattini *et al.* (PVLAS Collab.), Phys. Rev. Lett. **96**, 110406 (2006).
43. E. Zavattini *et al.* (PVLAS Collab.), Phys. Rev. **D77**, 032006 (2008).
44. F. Della Valle *et al.* (PVLAS Collab.), Eur. Phys. J. **C76**, 24 (2016).
45. E. Fischbach and C. Talmadge, Nature **356**, 207 (1992).
46. J.E. Moody and F. Wilczek, Phys. Rev. **D30**, 130 (1984);  
A.N. Youdin *et al.*, Phys. Rev. Lett. **77**, 2170 (1996);  
Wei-Tou Ni *et al.*, Phys. Rev. Lett. **82**, 2439 (1999);  
D.F. Phillips *et al.*, Phys. Rev. **D63**, 111101 (2001);  
B.R. Heckel *et al.* (Eöt-Wash Collab.), Phys. Rev. Lett. **97**, 021603 (2006);  
S.A. Hoedl *et al.*, Phys. Rev. Lett. **106**, 041801 (2011).
47. G. Raffelt, Phys. Rev. **D86**, 015001 (2012).
48. A. Arvanitaki and A.A. Geraci, Phys. Rev. Lett. **113**, 161801 (2014).
49. M.S. Turner, Phys. Reports **197**, 67 (1990);  
G.G. Raffelt, Phys. Reports **198**, 1 (1990).
50. G.G. Raffelt, *Stars as Laboratories for Fundamental Physics*, (Univ. of Chicago Press, Chicago, 1996).
51. G.G. Raffelt, Lect. Notes Phys. **741**, 51 (2008).
52. S. Andriamonje *et al.* (CAST Collab.), JCAP **0704**, 010 (2007).
53. P. Gondolo and G. Raffelt, Phys. Rev. **D79**, 107301 (2009).
54. H. Schlattl, A. Weiss, and G. Raffelt, Astropart. Phys. **10**, 353 (1999).
55. N. Vinyoles *et al.*, JCAP **1510**, 015 (2015).
56. A. Ayala *et al.*, Phys. Rev. Lett. **113**, 191302 (2014).
57. J. Redondo, JCAP **1312**, 008 (2013).
58. N. Viaux *et al.*, Phys. Rev. Lett. **111**, 231301 (2013).
59. M. Giannotti *et al.*, arXiv:1512.08108.
60. G.G. Raffelt, Phys. Lett. **B166**, 402 (1986);  
S.I. Blinnikov and N.V. Dunina-Barkovskaya, Mon. Not. R. Astron. Soc. **266**, 289 (1994).
61. M.M. Miller Bertolami *et al.*, JCAP **1410**, 069 (2014).
62. J. Isern *et al.*, Astrophys. J. Lett. **682**, L109 (2008);  
J. Isern *et al.*, J. Phys. Conf. Ser. **172**, 012005 (2009).
63. J. Isern *et al.*, Astron. & Astrophys. **512**, A86 (2010);  
A.H. Córscico *et al.*, Mon. Not. Roy. Astron. Soc. **424**, 2792 (2012);  
A.H. Córscico *et al.*, JCAP **1212**, 010 (2012).
64. J. Engel, D. Seckel, and A.C. Hayes, Phys. Rev. Lett. **65**, 960 (1990).
65. T. Moroi and H. Murayama, Phys. Lett. **B440**, 69 (1998).



See key on page 601

## Gauge & Higgs Boson Particle Listings

### Axions ( $A^0$ ) and Other Very Light Bosons

66. L.B. Leinson, JCAP **1408**, 031 (2014).
67. J. Keller and A. Sedrakian, Nucl. Phys. **A897**, 62 (2013); A. Sedrakian, [arXiv:1512.07828](#).
68. L.B. Leinson, Phys. Lett. **B741**, 87 (2015).
69. G.G. Raffelt, J. Redondo, and N. Viaux Maira, Phys. Rev. **D84**, 103008 (2011).
70. K. van Bibber *et al.*, Phys. Rev. **D39**, 2089 (1989).
71. D. Lazarus *et al.*, Phys. Rev. Lett. **69**, 2333 (1992).
72. S. Moriyama *et al.*, Phys. Lett. **B434**, 147 (1998); Y. Inoue *et al.*, Phys. Lett. **B536**, 18 (2002).
73. M. Minowa *et al.*, Phys. Lett. **B668**, 93 (2008).
74. E. Arik *et al.* (CAST Collab.), JCAP **0902**, 008 (2009).
75. S. Aune *et al.* (CAST Collab.), Phys. Rev. Lett. **107**, 261302 (2011); M. Arik *et al.* (CAST Collab.), Phys. Rev. Lett. **112**, 091302 (2014); M. Arik *et al.* (CAST Collab.), Phys. Rev. **D92**, 021101 (2015).
76. E. Aprile *et al.* (XENON100 Collab.), Phys. Rev. **D90**, 062009 (2014).
77. E. Armengaud *et al.*, JINST **9**, T05002 (2014).
78. K. Barth *et al.*, JCAP **1305**, 010 (2013).
79. F.T. Avignone III *et al.*, Phys. Rev. Lett. **81**, 5068 (1998); S. Cebrian *et al.*, Astropart. Phys. **10**, 397 (1999); A. Morales *et al.* (COSME Collab.), Astropart. Phys. **16**, 325 (2002); R. Bernabei *et al.*, Phys. Lett. **B515**, 6 (2001); Z. Ahmed *et al.* (CDMS Collab.), Phys. Rev. Lett. **103**, 141802 (2009); E. Armengaud *et al.* (EDELWEISS Collab.), JCAP **1311**, 067 (2013).
80. H. Davoudiasl and P. Huber, Phys. Rev. Lett. **97**, 141302 (2006).
81. H.S. Hudson *et al.*, ASP Conf. Ser. **455**, 25 (2012).
82. J.W. Brockway, E.D. Carlson, and G.G. Raffelt, Phys. Lett. **B383**, 439 (1996); J.A. Grifols, E. Massó, and R. Toldrà, Phys. Rev. Lett. **77**, 2372 (1996).
83. A. Payez *et al.*, JCAP **1502**, 006 (2015).
84. C. Csaki, N. Kaloper, and J. Terning, Phys. Rev. Lett. **88**, 161302 (2002).
85. A. Mirizzi, G.G. Raffelt, and P.D. Serpico, Lect. Notes Phys. **741**, 115 (2008).
86. R. Gill and J. S. Heyl, Phys. Rev. **D84**, 085001 (2011).
87. D. Horns *et al.*, Phys. Rev. **D85**, 085021 (2012).
88. A. Payez, J.R. Cudell, and D. Hutsemekers, JCAP **1207**, 041 (2012).
89. D. Horns and M. Meyer, JCAP **1202**, 033 (2012).
90. J. Biteau and D.A. Williams, Astrophys. J. **812**, 60 (2015).
91. A. De Angelis, G. Galanti, and M. Roncadelli, Phys. Rev. **D84**, 105030 (2011); M. Simet, D. Hooper, and P.D. Serpico, Phys. Rev. **D77**, 063001 (2008); M.A. Sanchez-Conde *et al.*, Phys. Rev. **D79**, 123511 (2009).
92. M. Meyer, D. Horns, and M. Raue, Phys. Rev. **D87**, 035027 (2013).
93. A. Abramowski *et al.* (H.E.S.S. Collab.), Phys. Rev. **D88**, 102003 (2013).
94. J.P. Conlon and M.C.D. Marsh, Phys. Rev. Lett. **111**, 151301 (2013); S. Angus *et al.*, JCAP **1409**, 026 (2014); D. Kraljic, M. Rummel, and J. P. Conlon, JCAP **1501**, 011 (2015).
95. A. Arvanitaki *et al.*, Phys. Rev. **D81**, 123530 (2010); A. Arvanitaki and S. Dubovsky, Phys. Rev. **D83**, 044026 (2011); A. Arvanitaki, M. Baryakhtar, and X. Huang, Phys. Rev. **D91**, 084011 (2015).
96. M.S. Turner, Phys. Rev. Lett. **59**, 2489 (1987) and Erratum *ibid.*, **60**, 1101 (1988); E. Massó, F. Rota, and G. Zsembinszki, Phys. Rev. **D66**, 023004 (2002); P. Graf and F. D. Steffen, Phys. Rev. **D83**, 075011 (2011).
97. S. Hannestad *et al.*, JCAP **1008**, 001 (2010); M. Archidiacono *et al.*, JCAP **1310**, 020 (2013); E. Di Valentino *et al.*, Phys. Lett. **B752**, 182 (2016).
98. M. Archidiacono *et al.*, JCAP **1505**, 050 (2015).
99. E. Massó and R. Toldrà, Phys. Rev. **D55**, 7967 (1997); D. Cadamuro and J. Redondo, JCAP **1202**, 032 (2012).
100. J. Preskill, M.B. Wise, and F. Wilczek, Phys. Lett. **B120**, 127 (1983); L.F. Abbott and P. Sikivie, Phys. Lett. **B120**, 133 (1983); M. Dine and W. Fischler, Phys. Lett. **B120**, 137 (1983).
101. K.J. Bae, J.-H. Huh, and J.E. Kim, JCAP **0809**, 005 (2008).
102. O. Wantz and E.P.S. Shellard, Phys. Rev. **D82**, 123508 (2010).
103. M. Tegmark *et al.*, Phys. Rev. **D73**, 023505 (2006).
104. M. Beltrán, J. García-Bellido, and J. Lesgourgues, Phys. Rev. **D75**, 103507 (2007); M.P. Hertzberg, M. Tegmark, and F. Wilczek, Phys. Rev. **D78**, 083507 (2008); J. Hamann *et al.*, JCAP **0906**, 022 (2009); P.A.R. Ade *et al.* [Planck Collab.], Astron. & Astrophys. **571**, A22 (2014); P.A.R. Ade *et al.* [Planck Collab.], Astrophys. Space Sci. **361**, 58 (2016).
105. P. Fox, A. Pierce, and S.D. Thomas, [hep-th/0409059](#); D.J.E. Marsh *et al.*, Phys. Rev. Lett. **113**, 011801 (2014); L. Visinelli and P. Gondolo, Phys. Rev. Lett. **113**, 011802 (2014).
106. P. Sikivie, Lect. Notes Phys. **741**, 19 (2008).
107. T. Hiramatsu *et al.*, Phys. Rev. **D83**, 123531 (2011); T. Hiramatsu *et al.*, Phys. Rev. **D85**, 105020 (2012) and Erratum *ibid.*, **86**, 089902 (2012); M. Kawasaki, K. Saikawa, and T. Sekiguchi, Phys. Rev. **D91**, 065014 (2015).
108. A. Ringwald and K. Saikawa, [arXiv:1512.06436](#).
109. E.W. Kolb and I.I. Tkachev, Phys. Rev. Lett. **71**, 3051 (1993), Astrophys. J. **460**, L25 (1996); K.M. Zurek, C.J. Hogan, and T.R. Quinn, Phys. Rev. **D75**, 043511 (2007).
110. E. Berkowitz, M.I. Buchoff, and E. Rinaldi, Phys. Rev. **D92**, 034507 (2015).
111. S. Borsanyi *et al.*, Phys. Lett. **B752**, 175 (2016); R. Kitano and N. Yamada, JHEP **1510**, 136 (2015).

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

112. B. Alles, M. D'Elia, and A. Di Giacomo, Phys. Lett. **B483**, 139 (2000);  
V. G. Bornyakov *et al.*, Phys. Rev. **D87**, 114508 (2013);  
C. Bonati *et al.*, arXiv:1512.06746 [hep-lat].
113. K.J. Bae, H. Baer, and A. Lessa, JCAP **1304**, 041 (2013);  
K.J. Bae, H. Baer, and E.J. Chun, Phys. Rev. **D89**, 03170 (2014);  
K.J. Bae, H. Baer, and E.J. Chun, JCAP **1312**, 028 (2013).
114. P. Arias *et al.*, JCAP **1206**, 013 (2012).
115. M. Bershadsky *et al.*, Phys. Rev. Lett. **66**, 1398 (1991);  
M. Ressel, Phys. Rev. **D44**, 3001 (1991).
116. D. Grin *et al.*, Phys. Rev. **D75**, 105018 (2007).
117. B.D. Blout *et al.*, Astrophys. J. **546**, 825 (2001).
118. P. Sikivie, Phys. Rev. **D32**, 2988 (1985);  
L. Krauss *et al.*, Phys. Rev. Lett. **55**, 1797 (1985);  
R. Bradley *et al.*, Rev. Mod. Phys. **75**, 777 (2003).
119. P. Sikivie and J. Ipser, Phys. Lett. **B291**, 288 (1992);  
P. Sikivie *et al.*, Phys. Rev. Lett. **75**, 2911 (1995).
120. S. DePanfilis *et al.*, Phys. Rev. Lett. **59**, 839 (1987);  
W. Wuensch *et al.*, Phys. Rev. **D40**, 3153 (1989);  
C. Hagmann *et al.*, Phys. Rev. **D42**, 1297 (1990).
121. S. Asztalos *et al.*, Phys. Rev. **D69**, 011101 (2004).
122. L. Duffy *et al.*, Phys. Rev. Lett. **95**, 091304 (2005);  
J. Hoskins *et al.*, Phys. Rev. **D84**, 121302 (2011).
123. S.J. Asztalos *et al.* (ADMX Collab.), Phys. Rev. Lett. **104**, 041301 (2010).
124. S.J. Asztalos *et al.*, Nucl. Instrum. Methods **A656**, 39 (2011).
125. S.J. Asztalos *et al.*, Phys. Rev. Lett. **104**, 041301 (2010).
126. G. Rybka *et al.*, Phys. Rev. Lett. **105**, 051801 (2010);  
A. Wagner *et al.*, Phys. Rev. Lett. **105**, 171801 (2010).
127. I. Ogawa, S. Matsuki, and K. Yamamoto, Phys. Rev. **D53**, 1740 (1996);  
Y. Kishimoto *et al.*, Phys. Lett. **A303**, 279 (2002);  
M. Tada *et al.*, Phys. Lett. **A303**, 285 (2002);  
T. Haseyama *et al.*, J. Low Temp. Phys. **150**, 549 (2008).
128. D. Horns *et al.*, JCAP **1304**, 016 (2013).
129. J. Suzuki *et al.*, JCAP **1509**, 042 (2015);  
B. Döbrich *et al.*, arXiv:1510.05869 [physics.ins-det].
130. G. Rybka *et al.*, Phys. Rev. **D91**, 011701 (2015).
131. P. Sikivie, Phys. Rev. Lett. **113**, 201301 (2014).
132. D. Budker *et al.*, Phys. Rev. **X4**, 021030 (2014).
133. P. Sikivie, N. Sullivan, and D. B. Tanner, Phys. Rev. Lett. **112**, 131301 (2014).
134. L. Krauss *et al.*, Phys. Rev. Lett. **55**, 1797 (1985);  
R. Barbieri *et al.*, Phys. Rev. **B226**, 357 (1989).
135. G. Ruoso *et al.*, arXiv:1511.09461 [hep-ph].

### $A^0$ (Axion) MASS LIMITS from Astrophysics and Cosmology

These bounds depend on model-dependent assumptions (i.e. — on a combination of axion parameters).

VALUE (MeV)	DOCUMENT ID	TECN.	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
>0.2	BARROSO	82	ASTR Standard Axion
>0.25	1 RAFFELT	82	ASTR Standard Axion
>0.2	2 DICUS	78c	ASTR Standard Axion
	MIKAELIAN	78	ASTR Stellar emission
>0.3	2 SATO	78	ASTR Standard Axion
>0.2	VYSOTSKII	78	ASTR Standard Axion

<sup>1</sup> Lower bound from 5.5 MeV  $\gamma$ -ray line from the sun.

<sup>2</sup> Lower bound from requiring the red giants' stellar evolution not be disrupted by axion emission.

### $A^0$ (Axion) and Other Light Boson ( $X^0$ ) Searches in Hadron Decays

Limits are for branching ratios.

VALUE	CL%	DOCUMENT ID	TECN.	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1 $\times 10^{-9}$	95	1 AAIJ	15AZ LHCb	$B^0 \rightarrow K^{*0} X^0 (X^0 \rightarrow \mu^+ \mu^-)$
<1.5 $\times 10^{-6}$	90	2 ADLARSON	13 WASA	$\pi^0 \rightarrow \gamma X^0 (X^0 \rightarrow e^+ e^-),$ $m_{X^0} = 100$ MeV
<2 $\times 10^{-8}$	90	3 BABUSCI	13B KLOE	$\phi \rightarrow \eta X^0 (X^0 \rightarrow e^+ e^-)$
		4 ARCHILLI	12 KLOE	$\phi \rightarrow \eta X^0, X^0 \rightarrow e^+ e^-$
<2 $\times 10^{-15}$	90	5 GNINENKO	12A BDMP	$\pi^0 \rightarrow \gamma X^0 (X^0 \rightarrow e^+ e^-)$
<3 $\times 10^{-14}$	90	6 GNINENKO	12B BDMP	$\eta(\eta') \rightarrow \gamma X^0 (X^0 \rightarrow e^+ e^-)$
<7 $\times 10^{-10}$	90	7 ADLER	04 B787	$K^+ \rightarrow \pi^+ X^0$
<7.3 $\times 10^{-11}$	90	8 ANISIMOVSK.	04 B949	$K^+ \rightarrow \pi^+ X^0$
<4.5 $\times 10^{-11}$	90	9 ADLER	02c B787	$K^+ \rightarrow \pi^+ X^0$
<4 $\times 10^{-5}$	90	10 ADLER	01 B787	$K^+ \rightarrow \pi^+ \pi^0 A^0$
<4.9 $\times 10^{-5}$	90	AMMAR	01B CLEO	$B^\pm \rightarrow \pi^\pm (K^\pm) X^0$
<5.3 $\times 10^{-5}$	90	AMMAR	01B CLEO	$B^0 \rightarrow K_L^0 X^0$
<3.3 $\times 10^{-5}$	90	11 ALTEGOER	98 NOMD	$\pi^0 \rightarrow \gamma X^0, m_{X^0} < 120$ MeV
<5.0 $\times 10^{-8}$	90	12 KITCHING	97 B787	$K^+ \rightarrow \pi^+ X^0 (X^0 \rightarrow \gamma \gamma)$
<5.2 $\times 10^{-10}$	90	13 ADLER	96 B787	$K^+ \rightarrow \pi^+ X^0$
<2.8 $\times 10^{-4}$	90	14 AMSLER	96B CBAR	$\pi^0 \rightarrow \gamma X^0, m_{X^0} < 65$ MeV
<3 $\times 10^{-4}$	90	14 AMSLER	96B CBAR	$\eta \rightarrow \gamma X^0, m_{X^0} = 50-200$ MeV
<4 $\times 10^{-5}$	90	14 AMSLER	96B CBAR	$\eta' \rightarrow \gamma X^0, m_{X^0} = 50-925$ MeV
<6 $\times 10^{-5}$	90	14 AMSLER	94B CBAR	$\pi^0 \rightarrow \gamma X^0, m_{X^0} = 65-125$ MeV
<6 $\times 10^{-5}$	90	14 AMSLER	94B CBAR	$\eta \rightarrow \gamma X^0, m_{X^0} = 200-525$ MeV
<7 $\times 10^{-3}$	90	15 MEIJERDREES	94 CNTR	$\pi^0 \rightarrow \gamma X^0, m_{X^0} = 25$ MeV
<2 $\times 10^{-3}$	90	15 MEIJERDREES	94 CNTR	$\pi^0 \rightarrow \gamma X^0, m_{X^0} = 100$ MeV
<2 $\times 10^{-7}$	90	16 ATIYA	93B B787	Sup. by ADLER 04
<3 $\times 10^{-13}$	90	17 NG	93 COSM	$\pi^0 \rightarrow \gamma X^0$
<1.1 $\times 10^{-8}$	90	18 ALLIEGRO	92 SPEC	$K^+ \rightarrow \pi^+ X^0 (X^0 \rightarrow e^+ e^-)$
<5 $\times 10^{-4}$	90	19 ATIYA	92 B787	$\pi^0 \rightarrow \gamma X^0$
<1 $\times 10^{-12}$	95	20 BARABASH	92 BDMP	$\pi^\pm \rightarrow e^\pm \nu X^0 (X^0 \rightarrow e^+ e^-, \gamma \gamma),$ $m_{X^0} = 8$ MeV
<1 $\times 10^{-12}$	95	21 BARABASH	92 BDMP	$K^\pm \rightarrow \pi^\pm X^0 (X^0 \rightarrow e^+ e^-, \gamma \gamma),$ $m_{X^0} = 10$ MeV
<1 $\times 10^{-11}$	95	22 BARABASH	92 BDMP	$K_L^0 \rightarrow \pi^0 X^0 (X^0 \rightarrow e^+ e^-, \gamma \gamma),$ $m_{X^0} = 10$ MeV
<1 $\times 10^{-14}$	95	23 BARABASH	92 BDMP	$\eta' \rightarrow \eta X^0 (X^0 \rightarrow e^+ e^-, \gamma \gamma),$ $m_{X^0} = 10$ MeV
<4 $\times 10^{-6}$	90	24 MEIJERDREES	92 SPEC	$\pi^0 \rightarrow \gamma X^0 (X^0 \rightarrow e^+ e^-),$ $m_{X^0} = 100$ MeV
<1 $\times 10^{-7}$	90	25 ATIYA	90B B787	Sup. by KITCHING 97
<1.3 $\times 10^{-8}$	90	26 KORENCHEN...	87 SPEC	$\pi^+ \rightarrow e^+ \nu A^0 (A^0 \rightarrow e^+ e^-)$
<1 $\times 10^{-9}$	90	27 EICHLER	86 SPEC	Stopped $\pi^+ \rightarrow e^+ \nu A^0$
<2 $\times 10^{-5}$	90	28 YAMAZAKI	84 SPEC	For $160 < m < 260$ MeV
<(1.5-4) $\times 10^{-6}$	90	28 YAMAZAKI	84 SPEC	$K$ decay, $m_{X^0} \ll 100$ MeV
		29 ASANO	82 CNTR	Stopped $K^+ \rightarrow \pi^+ X^0$
		30 ASANO	81B CNTR	Stopped $K^+ \rightarrow \pi^+ X^0$
		31 ZHITNITSKII	79	Heavy axion

<sup>1</sup> The limit is for  $\tau_{X^0} = 10$  ps and  $m_{X^0} = 214-4350$  MeV. See their Fig. 4 for mass- and lifetime-dependent limits.

<sup>2</sup> Limits between  $2.0 \times 10^{-5}$  and  $1.5 \times 10^{-6}$  are obtained for  $m_{X^0} = 20-100$  MeV (see their Fig. 8). Angular momentum conservation requires that  $X^0$  has spin  $\geq 1$ .

<sup>3</sup> The limit is for  $B(\phi \rightarrow \eta X^0) \cdot B(X^0 \rightarrow e^+ e^-)$  and applies to  $m_{X^0} = 410$  MeV. It is derived by analyzing  $\eta \rightarrow \pi^0 \pi^0 \pi^0$  and  $\pi^- \pi^+ \pi^0$ . Limits between  $1 \times 10^{-6}$  and  $2 \times 10^{-8}$  are obtained for  $m_{X^0} \leq 450$  MeV (see their Fig. 6).

<sup>4</sup> ARCHILLI 12 analyzed  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decays. Derived limits on  $\alpha'/\alpha < 2 \times 10^{-5}$  for  $m_{X^0} = 50-420$  MeV at 90% CL. See their Fig. 8 for mass-dependent limits.

<sup>5</sup> This limit is for  $B(\pi^0 \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow e^+ e^-)$  and applies for  $m_{X^0} = 90$  MeV and  $\tau_{X^0} \simeq 1 \times 10^{-8}$  sec. Limits between  $10^{-8}$  and  $2 \times 10^{-15}$  are obtained for  $m_{X^0} = 3-120$  MeV and  $\tau_{X^0} = 1 \times 10^{-11}-1$  sec. See their Fig. 3 for limits at different masses and lifetimes.

<sup>6</sup> This limit is for  $B(\eta \rightarrow \gamma X^0) \cdot B(X^0 \rightarrow e^+ e^-)$  and applies for  $m_{X^0} = 100$  MeV and  $\tau_{X^0} \simeq 6 \times 10^{-9}$  sec. Limits between  $10^{-5}$  and  $3 \times 10^{-14}$  are obtained for  $m_{X^0} \lesssim 550$  MeV and  $\tau_{X^0} = 10^{-10}-10$  sec. See their Fig. 5 for limits at different mass and lifetime and for  $\eta'$  decays.

<sup>7</sup> This limit applies for a mass near 180 MeV. For other masses in the range  $m_{X^0} = 150-250$  MeV the limit is less restrictive, but still improves ADLER 02c and ATIYA 93B.

<sup>8</sup> ANISIMOVSKY 04 bound is for  $m_{X^0} = 0$ .

<sup>9</sup> ADLER 02c bound is for  $m_{X^0} < 60$  MeV. See Fig. 2 for limits at higher masses.

<sup>10</sup> The quoted limit is for  $m_{X^0} = 0-80$  MeV. See their Fig. 5 for the limit at higher mass. The branching fraction limit assumes pure phase space decay distributions.

<sup>11</sup> ALTEGOER 98 looked for  $X^0$  from  $\pi^0$  decay which penetrate the shielding and convert to  $\pi^0$  in the external Coulomb field of a nucleus.

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# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

- <sup>12</sup> KITCHING 97 limit is for  $B(K^+ \rightarrow \pi^+ X^0) \cdot B(X^0 \rightarrow \gamma\gamma)$  and applies for  $m_{X^0} \simeq 50$  MeV,  $\tau_{X^0} < 10^{-10}$  s. Limits are provided for  $0 < m_{X^0} < 100$  MeV,  $\tau_{X^0} < 10^{-8}$  s.
- <sup>13</sup> ADLER 96 looked for a peak in missing-mass distribution. This work is an update of ATIYA 93. The limit is for massless stable  $X^0$  particles and extends to  $m_{X^0}=80$  MeV at the same level. See paper for dependence on finite lifetime.
- <sup>14</sup> AMSLER 94B and AMSLER 96B looked for a peak in missing-mass distribution.
- <sup>15</sup> The MEIJERDREES 94 limit is based on inclusive photon spectrum and is independent of  $X^0$  decay modes. It applies to  $\tau(X^0) > 10^{-23}$  sec.
- <sup>16</sup> ATIYA 93B looked for a peak in missing mass distribution. The bound applies for stable  $X^0$  of  $m_{X^0}=150$ –250 MeV, and the limit becomes stronger ( $10^{-8}$ ) for  $m_{X^0}=180$ –240 MeV.
- <sup>17</sup> NG 93 studied the production of  $X^0$  via  $\gamma\gamma \rightarrow \pi^0 \rightarrow \gamma X^0$  in the early universe at  $T \simeq 1$  MeV. The bound on extra neutrinos from nucleosynthesis  $\Delta N_\nu < 0.3$  (WALKER 91) is employed. It applies to  $m_{X^0} \ll 1$  MeV in order to be relativistic down to nucleosynthesis temperature. See paper for heavier  $X^0$ .
- <sup>18</sup> ALLIEGRO 92 limit applies for  $m_{X^0}=150$ –340 MeV and is the branching ratio times the decay probability. Limit is  $< 1.5 \times 10^{-8}$  at 99%CL.
- <sup>19</sup> ATIYA 92 looked for a peak in missing mass distribution. The limit applies to  $m_{X^0}=0$ –130 MeV in the narrow resonance limit. See paper for the dependence on lifetime. Covariance requires  $X^0$  to be a vector particle.
- <sup>20</sup> BARABASH 92 is a beam dump experiment that searched for a light Higgs. Limits between  $1 \times 10^{-12}$  and  $1 \times 10^{-7}$  are obtained for  $3 < m_{X^0} < 40$  MeV.
- <sup>21</sup> Limits between  $1 \times 10^{-12}$  and 1 are obtained for  $4 < m_{X^0} < 69$  MeV.
- <sup>22</sup> Limits between  $1 \times 10^{-11}$  and  $5 \times 10^{-3}$  are obtained for  $4 < m_{X^0} < 63$  MeV.
- <sup>23</sup> Limits between  $1 \times 10^{-14}$  and 1 are obtained for  $3 < m_{X^0} < 82$  MeV.
- <sup>24</sup> MEIJERDREES 92 limit applies for  $\tau_{X^0} = 10^{-23}$ – $10^{-11}$  sec. Limits between  $2 \times 10^{-4}$  and  $4 \times 10^{-6}$  are obtained for  $m_{X^0} = 25$ –120 MeV. Angular momentum conservation requires that  $X^0$  has spin  $\geq 1$ .
- <sup>25</sup> ATIYA 90B limit is for  $B(K^+ \rightarrow \pi^+ X^0) \cdot B(X^0 \rightarrow \gamma\gamma)$  and applies for  $m_{X^0} = 50$  MeV,  $\tau_{X^0} < 10^{-10}$  s. Limits are also provided for  $0 < m_{X^0} < 100$  MeV,  $\tau_{X^0} < 10^{-8}$  s.
- <sup>26</sup> KORENCHENKO 87 limit assumes  $m_{A^0} = 1.7$  MeV,  $\tau_{A^0} \lesssim 10^{-12}$  s, and  $B(A^0 \rightarrow e^+ e^-) = 1$ .
- <sup>27</sup> EICHLER 86 looked for  $\pi^+ \rightarrow e^+ \nu A^0$  followed by  $A^0 \rightarrow e^+ e^-$ . Limits on the branching fraction depend on the mass and and lifetime of  $A^0$ . The quoted limits are valid when  $\tau(A^0) \gtrsim 3 \cdot 10^{-10}$  s if the decays are kinematically allowed.
- <sup>28</sup> YAMAZAKI 84 looked for a discrete line in  $K^+ \rightarrow \pi^+ X$ . Sensitive to wide mass range (5–300 MeV), independent of whether X decays promptly or not.
- <sup>29</sup> ASANO 82 at KEK set limits for  $B(K^+ \rightarrow \pi^+ X^0)$  for  $m_{X^0} < 100$  MeV as  $BR < 4 \cdot 10^{-8}$  for  $\tau(X^0 \rightarrow n\gamma\text{'s}) > 1 \cdot 10^{-9}$  s,  $BR < 1.4 \times 10^{-6}$  for  $\tau < 1 \cdot 10^{-9}$  s.
- <sup>30</sup> ASANO 81B is KEK experiment. Set  $B(K^+ \rightarrow \pi^+ X^0) < 3.8 \times 10^{-8}$  at CL = 90%.
- <sup>31</sup> ZHITNITSKII 79 argue that a heavy axion predicted by YANG 78 ( $3 < m < 40$  MeV) contradicts experimental muon anomalous magnetic moments.

### $A^0$ (Axion) Searches in Quarkonium Decays

Decay or transition of quarkonium. Limits are for branching ratio.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.0 \times 10^{-5}$	90	<sup>1</sup> ANTREASYAN 90C	CBAL	$T(1S) \rightarrow A^0 \gamma$
$< 5 \times 10^{-5}$	90	<sup>2</sup> DRUZHININ 87	ND	$\phi \rightarrow A^0 \gamma (A^0 \rightarrow e^+ e^-)$
$< 2 \times 10^{-3}$	90	<sup>3</sup> DRUZHININ 87	ND	$\phi \rightarrow A^0 \gamma (A^0 \rightarrow \gamma\gamma)$
$< 7 \times 10^{-6}$	90	<sup>4</sup> DRUZHININ 87	ND	$\phi \rightarrow A^0 \gamma (A^0 \rightarrow \text{missing})$
$< 1.4 \times 10^{-5}$	90	<sup>5</sup> EDWARDS 82	CBAL	$J/\psi \rightarrow A^0 \gamma$
<sup>1</sup> ANTREASYAN 90C assume that $A^0$ does not decay in the detector.				
<sup>2</sup> The first DRUZHININ 87 limit is valid when $\tau_{A^0}/m_{A^0} < 3 \times 10^{-13}$ s/MeV and $m_{A^0} < 20$ MeV.				
<sup>3</sup> The second DRUZHININ 87 limit is valid when $\tau_{A^0}/m_{A^0} < 5 \times 10^{-13}$ s/MeV and $m_{A^0} < 20$ MeV.				
<sup>4</sup> The third DRUZHININ 87 limit is valid when $\tau_{A^0}/m_{A^0} > 7 \times 10^{-12}$ s/MeV and $m_{A^0} < 200$ MeV.				
<sup>5</sup> EDWARDS 82 looked for $J/\psi \rightarrow \gamma A^0$ decays by looking for events with a single $\gamma$ [of energy $\sim 1/2$ the $J/\psi(1S)$ mass], plus nothing else in the detector. The limit is inconsistent with the axion interpretation of the FAISSNER 81B result.				

### $A^0$ (Axion) Searches in Positronium Decays

Decay or transition of positronium. Limits are for branching ratio.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.4 \times 10^{-5}$	90	<sup>1</sup> BADERT... 02	CNTR	$\alpha\text{-Ps} \rightarrow \gamma X_1 X_2, m_{X_1} + m_{X_2} \leq 900 \text{ keV}$
$< 2 \times 10^{-4}$	90	MAENO 95	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma, m_{A^0}=850\text{--}1013 \text{ keV}$
$< 3.0 \times 10^{-4}$	90	<sup>2</sup> ASAI 94	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma, m_{A^0}=30\text{--}500 \text{ keV}$
$< 2.8 \times 10^{-5}$	90	<sup>3</sup> AKOPYAN 91	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma (A^0 \rightarrow \gamma\gamma), m_{A^0} < 30 \text{ keV}$
$< 1.1 \times 10^{-6}$	90	<sup>4</sup> ASAI 91	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma, m_{A^0} < 800 \text{ keV}$
$< 3.8 \times 10^{-4}$	90	GNINENKO 90	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma, m_{A^0} < 30 \text{ keV}$
$< (1.5) \times 10^{-4}$	95	<sup>5</sup> TSUCHIAKI 90	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma, m_{A^0} = 300\text{--}900 \text{ keV}$
$< 6.4 \times 10^{-5}$	90	<sup>6</sup> ORITO 89	CNTR	$\alpha\text{-Ps} \rightarrow A^0 \gamma, m_{A^0} < 30 \text{ keV}$
		<sup>7</sup> AMALDI 85	CNTR	Ortho-positronium
		<sup>8</sup> CARBONI 83	CNTR	Ortho-positronium

- <sup>1</sup> BADERTSCHER 02 looked for a three-body decay of ortho-positronium into a photon and two penetrating (neutral or milli-charged) particles.
- <sup>2</sup> The ASAI 94 limit is based on inclusive photon spectrum and is independent of  $A^0$  decay modes.
- <sup>3</sup> The AKOPYAN 91 limit applies for a short-lived  $A^0$  with  $\tau_{A^0} < 10^{-13}$   $m_{A^0}$  [keV] s.
- <sup>4</sup> ASAI 91 limit translates to  $g_{A^0 e^+ e^-}^2 / 4\pi < 1.1 \times 10^{-11}$  (90% CL) for  $m_{A^0} < 800$  keV.
- <sup>5</sup> The TSUCHIAKI 90 limit is based on inclusive photon spectrum and is independent of  $A^0$  decay modes.
- <sup>6</sup> ORITO 89 limit translates to  $g_{A^0 e^+ e^-}^2 / 4\pi < 6.2 \times 10^{-10}$ . Somewhat more sensitive limits are obtained for larger  $m_{A^0}$ :  $B < 7.6 \times 10^{-6}$  at 100 keV.
- <sup>7</sup> AMALDI 85 set limits  $B(A^0 \gamma) / B(\gamma\gamma\gamma) < (1\text{--}5) \times 10^{-6}$  for  $m_{A^0} = 900\text{--}100 \text{ keV}$  which are about 1/10 of the CARBONI 83 limits.
- <sup>8</sup> CARBONI 83 looked for orthopositronium  $\rightarrow A^0 \gamma$ . Set limit for  $A^0$  electron coupling squared,  $g(e e A^0)^2 / (4\pi) < 6 \cdot 10^{-10}$ – $7 \cdot 10^{-9}$  for  $m_{A^0}$  from 150–900 keV (CL = 99.7%). This is about 1/10 of the bound from  $g\text{--}2$  experiments.

### $A^0$ (Axion) Search in Photoproduction

VALUE	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<sup>1</sup> BASSOMPIERRE... 95	$m_{A^0} = 1.8 \pm 0.2 \text{ MeV}$	
<sup>1</sup> BASSOMPIERRE 95 is an extension of BASSOMPIERRE 93. They looked for a peak in the invariant mass of $e^+ e^-$ pairs in the region $m_{e^+ e^-} = 1.8 \pm 0.2 \text{ MeV}$ . They obtained bounds on the production rate $A^0$ for $\tau(A^0) = 10^{-18}$ – $10^{-9}$ sec. They also found an excess of events in the range $m_{e^+ e^-} = 2.1\text{--}3.5 \text{ MeV}$ .		

### $A^0$ (Axion) Production in Hadron Collisions

Limits are for  $\sigma(A^0) / \sigma(\pi^0)$ .

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
			<sup>1</sup> JAIN	07	CNTR $A^0 \rightarrow e^+ e^-$
			<sup>2</sup> AHMAD	97	SPEC $e^+$ production
			<sup>3</sup> LEINBERGER	97	SPEC $A^0 \rightarrow e^+ e^-$
			<sup>4</sup> GANZ	96	SPEC $A^0 \rightarrow e^+ e^-$
			<sup>5</sup> KAMEL	96	EMUL $^{32}\text{S}$ emulsion, $A^0 \rightarrow e^+ e^-$
			<sup>6</sup> BLUEMLEIN	92	BDMP $A^0 N_Z \rightarrow \ell^+ \ell^- N_Z$
			<sup>7</sup> MEIJERDREES 92	SPEC	$\pi^- p \rightarrow n A^0, A^0 \rightarrow e^+ e^-$
			<sup>8</sup> BLUEMLEIN	91	BDMP $A^0 \rightarrow e^+ e^-, 2\gamma$
			<sup>9</sup> FAISSNER	89	OSPK Beam dump, $A^0 \rightarrow e^+ e^-$
			<sup>10</sup> DEBOER	88	RVUE $A^0 \rightarrow e^+ e^-$
			<sup>11</sup> EL-NADI	88	EMUL $A^0 \rightarrow e^+ e^-$
			<sup>12</sup> FAISSNER	88	OSPK Beam dump, $A^0 \rightarrow 2\gamma$
			<sup>13</sup> BADIER	86	BDMP $A^0 \rightarrow e^+ e^-$
$< 2 \cdot 10^{-11}$	90	0	<sup>14</sup> BERGSMA	85	CHRM CERN beam dump
$< 1 \cdot 10^{-13}$	90	0	<sup>14</sup> BERGSMA	85	CHRM CERN beam dump
		24	<sup>15</sup> FAISSNER	83	OSPK Beam dump, $A^0 \rightarrow 2\gamma$
			<sup>16</sup> FAISSNER	83B	RVUE LAMPF beam dump
			<sup>17</sup> FRANK	83B	RVUE LAMPF beam dump
			<sup>18</sup> HOFFMAN	83	CNTR $\pi p \rightarrow n A^0 (A^0 \rightarrow e^+ e^-)$
			<sup>19</sup> FETSCHER	82	RVUE See FAISSNER 81B
		12	<sup>20</sup> FAISSNER	81	OSPK CERN PS $\nu$ wideband
		15	<sup>21</sup> FAISSNER	81B	OSPK Beam dump, $A^0 \rightarrow 2\gamma$
		8	<sup>22</sup> KIM	81	OSPK 26 GeV $p N \rightarrow A^0 X$
		0	<sup>23</sup> FAISSNER	80	OSPK Beam dump, $A^0 \rightarrow e^+ e^-$
$< 1 \cdot 10^{-8}$	90		<sup>24</sup> JACQUES	80	HLBC 28 GeV protons
$< 1 \cdot 10^{-14}$	90		<sup>24</sup> JACQUES	80	HLBC Beam dump
			<sup>25</sup> SOUKAS	80	CALO 28 GeV $p$ beam dump
			<sup>26</sup> BECHIS	79	CNTR
$< 1 \cdot 10^{-8}$	90		<sup>27</sup> COTEUS	79	OSPK Beam dump
$< 1 \cdot 10^{-3}$	95		<sup>28</sup> DISHAW	79	CALO 400 GeV $pp$
$< 1 \cdot 10^{-8}$	90		ALIBRAN	78	HYBR Beam dump
$< 6 \cdot 10^{-9}$	95		ASRATYAN	78B	CALO Beam dump
$< 1.5 \times 10^{-8}$	90		<sup>29</sup> BELLOTTI	78	HLBC Beam dump
$< 5.4 \times 10^{-14}$	90		<sup>29</sup> BELLOTTI	78	HLBC $m_{A^0}=1.5 \text{ MeV}$
$< 4.1 \times 10^{-9}$	90		<sup>29</sup> BELLOTTI	78	HLBC $m_{A^0}=1 \text{ MeV}$
$< 1 \cdot 10^{-8}$	90		<sup>30</sup> BOSETTI	78B	HYBR Beam dump
			<sup>31</sup> DONNELLY	78	
$< 0.5 \times 10^{-8}$	90		HANSL	78D	WIRE Beam dump
			<sup>32</sup> MICELMAC...	78	
			<sup>33</sup> VYSOTSKII	78	

- <sup>1</sup> JAIN 07 claims evidence for  $A^0 \rightarrow e^+ e^-$  produced in  $^{207}\text{Pb}$  collision on nuclear emulsion (Ag/Br) for  $m(A^0) = 7 \pm 1$  or  $19 \pm 1 \text{ MeV}$  and  $\tau(A^0) \leq 10^{-13} \text{ s}$ .
- <sup>2</sup> AHMAD 97 reports a result of APEX Collaboration which studied positron production in  $^{238}\text{U} + ^{232}\text{Ta}$  and  $^{238}\text{U} + ^{181}\text{Ta}$  collisions, without requiring a coincident electron. No narrow lines were found for  $250 < E_{e^+} < 750 \text{ keV}$ .

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

- <sup>3</sup> LEINBERGER 97 (ORANGE Collaboration) at GSI looked for a narrow sum-energy  $e^+e^-$  line at  $\sim 635$  keV in  $^{238}\text{U}+^{181}\text{Ta}$  collision. Limits on the production probability for a narrow sum-energy  $e^+e^-$  line are set. See their Table 2.
- <sup>4</sup> GANZ 96 (EPOS II Collaboration) has placed upper bounds on the production cross section of  $e^+e^-$  pairs from  $^{238}\text{U}+^{181}\text{Ta}$  and  $^{238}\text{U}+^{232}\text{Th}$  collisions at GSI. See Table 2 for limits both for back-to-back and isotropic configurations of  $e^+e^-$  pairs. These limits rule out the existence of peaks in the  $e^+e^-$  sum-energy distribution, reported by an earlier version of this experiment.
- <sup>5</sup> KAMEL 96 looked for  $e^+e^-$  pairs from the collision of  $^{32}\text{S}$  (200 GeV/nucleon) and emulsion. No evidence of mass peaks is found in the region of sensitivity  $m_{ee} > 2$  MeV.
- <sup>6</sup> BLUEMLEIN 92 is a proton beam dump experiment at Serpukhov with a secondary target to induce Bethe-Heitler production of  $e^+e^-$  or  $\mu^+\mu^-$  from the produce  $A^0$ . See Fig. 5 for the excluded region in  $m_{A^0}$ - $x$  plane. For the standard axion,  $0.3 < x < 25$  is excluded at 95% CL. If combined with BLUEMLEIN 91,  $0.008 < x < 32$  is excluded.
- <sup>7</sup> MEIJERDREES 92 give  $\Gamma(\pi^-\rho \rightarrow nA^0)\cdot\text{B}(A^0 \rightarrow e^+e^-)/\Gamma(\pi^-\rho \rightarrow \text{all}) < 10^{-5}$  (90% CL) for  $m_{A^0} = 100$  MeV,  $\tau_{A^0} = 10^{-11}-10^{-23}$  sec. Limits ranging from  $2.5 \times 10^{-3}$  to  $10^{-7}$  are given for  $m_{A^0} = 25-136$  MeV.
- <sup>8</sup> BLUEMLEIN 91 is a proton beam dump experiment at Serpukhov. No candidate event for  $A^0 \rightarrow e^+e^-$ ,  $2\gamma$  are found. Fig. 6 gives the excluded region in  $m_{A^0}$ - $x$  plane ( $x = \tan\beta = v_2/v_1$ ). Standard axion is excluded for  $0.2 < m_{A^0} < 3.2$  MeV for most  $x > 1$ ,  $0.2-11$  MeV for most  $x < 1$ .
- <sup>9</sup> FAISSNER 89 searched for  $A^0 \rightarrow e^+e^-$  in a proton beam dump experiment at SIN. No evidence of events was observed over the background. A standard axion with mass  $2m_e-20$  MeV is excluded. Lower limit on  $f_{A^0}$  of  $\simeq 10^4$  GeV is given for  $m_{A^0} = 2m_e-20$  MeV.
- <sup>10</sup> DEBOER 88 reanalyze EL-NADI 88 data and claim evidence for three distinct states with mass  $\sim 1.1$ ,  $\sim 2.1$ , and  $\sim 9$  MeV, lifetimes  $10^{-16}-10^{-15}$  s decaying to  $e^+e^-$  and note the similarity of the data with those of a cosmic-ray experiment by Bristol group (B.M. Anand, Proc. of the Royal Society of London, Section A **A22** 183 (1953)). For a criticism see PERKINS 89, who suggests that the events are compatible with  $\pi^0$  Dalitz decay. DEBOER 89b is a reply which contests the criticism.
- <sup>11</sup> EL-NADI 88 claim the existence of a neutral particle decaying into  $e^+e^-$  with mass  $1.60 \pm 0.59$  MeV, lifetime  $(0.15 \pm 0.01) \times 10^{-14}$  s, which is produced in heavy ion interactions with emulsion nuclei at  $\sim 4$  GeV/c/nucleon.
- <sup>12</sup> FAISSNER 88 is a proton beam dump experiment at SIN. They found no candidate event for  $A^0 \rightarrow \gamma\gamma$ . A standard axion decaying to  $2\gamma$  is excluded except for a region  $x \simeq 1$ . Lower limit on  $f_{A^0}$  of  $10^2-10^3$  GeV is given for  $m_{A^0} = 0.1-1$  MeV.
- <sup>13</sup> BADIER 86 did not find long-lived  $A^0$  in 300 GeV  $\pi^-$  Beam Dump Experiment that decays into  $e^+e^-$  in the mass range  $m_{A^0} = (20-200)$  MeV, which excludes the  $A^0$  decay constant  $f(A^0)$  in the interval (60-600) GeV. See their figure 6 for excluded region on  $f(A^0)$ - $m_{A^0}$  plane.
- <sup>14</sup> BERGSMAN 85 look for  $A^0 \rightarrow 2\gamma$ ,  $e^+e^-$ ,  $\mu^+\mu^-$ . First limit above is for  $m_{A^0} = 1$  MeV; second is for 200 MeV. See their figure 4 for excluded region on  $f_{A^0}$ - $m_{A^0}$  plane, where  $f_{A^0}$  is  $A^0$  decay constant. For Peccei-Quinn PECEI 77  $A^0$ ,  $m_{A^0} < 180$  keV and  $\tau > 0.037$  s. (CL = 90%). For the axion of FAISSNER 81b at 250 keV, BERGSMAN 85 expect 15 events but observe zero.
- <sup>15</sup> FAISSNER 83 observed 19  $1\gamma$  and 12  $2\gamma$  events where a background of 4.8 and 2.3 respectively is expected. A small-angle peak is observed even if iron wall is set in front of the decay region.
- <sup>16</sup> FAISSNER 83b extrapolate SIN  $\gamma$  signal to LAMPF  $\nu$  experimental condition. Resulting 370  $\gamma$ 's are not at variance with LAMPF upper limit of 450  $\gamma$ 's. Derived from LAMPF limit that  $[d\sigma(A^0)/d\omega \text{ at } 90^\circ] m_{A^0}/\tau_{A^0} < 14 \times 10^{-35} \text{ cm}^2 \text{ sr}^{-1} \text{ MeV ms}^{-1}$ . See comment on FRANK 83b.
- <sup>17</sup> FRANK 83b stress the importance of LAMPF data bins with negative net signal. By statistical analysis say that LAMPF and SIN- $A^0$  are at variance when extrapolation by phase-space model is done. They find LAMPF upper limit is 248 not 450  $\gamma$ 's. See comment on FAISSNER 83b.
- <sup>18</sup> HOFFMAN 83 set CL = 90% limit  $d\sigma/dt \text{ B}(e^+e^-) < 3.5 \times 10^{-32} \text{ cm}^2/\text{GeV}^2$  for  $140 < m_{A^0} < 160$  MeV. Limit assumes  $\tau(A^0) < 10^{-9}$  s.
- <sup>19</sup> FETSCHER 82 reanalyzes SIN beam-dump data of FAISSNER 81. Claims no evidence for axion since  $2\gamma$  peak rate remarkably decreases if iron wall is set in front of the decay region.
- <sup>20</sup> FAISSNER 81 see excess  $\mu e$  events. Suggest axion interactions.
- <sup>21</sup> FAISSNER 81b is SIN 590 MeV proton beam dump. Observed  $14.5 \pm 5.0$  events of  $2\gamma$  decay of long-lived neutral penetrating particle with  $m_{2\gamma} \lesssim 1$  MeV. Axion interpretation with  $\eta$ - $A^0$  mixing gives  $m_{A^0} = 250 \pm 25$  keV,  $\tau_{(2\gamma)} = (7.3 \pm 3.7) \times 10^{-3}$  s from above rate. See critical remarks below in comments of FETSCHER 82, FAISSNER 83, FAISSNER 83b, FRANK 83b, and BERGSMAN 85. Also see in the next subsection ALEKSEEV 82b, CAVAGNAC 83, and ANANEV 85.
- <sup>22</sup> KIM 81 analyzed 8 candidates for  $A^0 \rightarrow 2\gamma$  obtained by Aachen-Padova experiment at CERN with 26 GeV protons on Be. Estimated axion mass is about 300 keV and lifetime is  $(0.86 \sim 5.6) \times 10^{-3}$  s depending on models. Faissner (private communication), says axion production underestimated and mass overestimated. Correct value around 200 keV.
- <sup>23</sup> FAISSNER 80 is SIN beam dump experiment with 590 MeV protons looking for  $A^0 \rightarrow e^+e^-$  decay. Assuming  $A^0/\pi^0 = 5.5 \times 10^{-7}$ , obtained decay rate limit  $20/(A^0 \text{ mass})$  MeV/s (CL = 90%), which is about  $10^{-7}$  below theory and interpreted as upper limit to  $m_{A^0} < 2m_e$ .
- <sup>24</sup> JACQUES 80 is a BNL beam dump experiment. First limit above comes from nonobservation of excess neutral-current-type events  $[\sigma(\text{production})\sigma(\text{interaction}) < 7 \times 10^{-68} \text{ cm}^4]$ , CL = 90%. Second limit is from nonobservation of axion decays into  $2\gamma$ 's or  $e^+e^-$ , and for axion mass a few MeV.
- <sup>25</sup> SOUKAS 80 at BNL observed no excess of neutral-current-type events in beam dump.
- <sup>26</sup> BECHIS 79 looked for the axion production in low energy electron Bremsstrahlung and the subsequent decay into either  $2\gamma$  or  $e^+e^-$ . No signal found. CL = 90% limits for model parameter(s) are given.
- <sup>27</sup> COTEUS 79 is a beam dump experiment at BNL.

- <sup>28</sup> DISHAW 79 is a calorimetric experiment and looks for low energy tail of energy distributions due to energy lost to weakly interacting particles.
- <sup>29</sup> BELLOTTI 78 first value comes from search for  $A^0 \rightarrow e^+e^-$ . Second value comes from search for  $A^0 \rightarrow 2\gamma$ , assuming mass  $< 2m_e$ . For any mass satisfying this, limit is above value  $\times (\text{mass}^{-4})$ . Third value uses data of PL 60B 401 and quotes  $\sigma(\text{production})\sigma(\text{interaction}) < 10^{-67} \text{ cm}^4$ .
- <sup>30</sup> BOSETTI 78B quotes  $\sigma(\text{production})\sigma(\text{interaction}) < 2 \times 10^{-67} \text{ cm}^4$ .
- <sup>31</sup> DONNELLY 78 examines data from reactor neutrino experiments of REINES 76 and GURR 74 as well as SLAC beam dump experiment. Evidence is negative.
- <sup>32</sup> MICELMACHER 78 finds no evidence of axion existence in reactor experiments of REINES 76 and GURR 74. (See reference under DONNELLY 78 below).
- <sup>33</sup> VYSOTSKI 78 derived lower limit for the axion mass 25 keV from luminosity of the sun and 200 keV from red supergiants.

### $A^0$ (Axion) Searches in Reactor Experiments

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1	CHANG 07		Primakoff or Compton
2	ALTMANN 95	CNTR	Reactor; $A^0 \rightarrow e^+e^-$
3	KETOV 86	SPEC	Reactor; $A^0 \rightarrow \gamma\gamma$
4	KOCH 86	SPEC	Reactor; $A^0 \rightarrow \gamma\gamma$
5	DATAR 82	CNTR	Light water reactor
6	VUILLEUMIER 81	CNTR	Reactor; $A^0 \rightarrow 2\gamma$

- <sup>1</sup> CHANG 07 looked for monochromatic photons from Primakoff or Compton conversion of axions from the Kuo-Sheng reactor due to axion coupling to photon or electron, respectively. The search places model-independent limits on the products  $G_{A\gamma\gamma} G_{ANN}$  and  $G_{Aee} G_{ANN}$  for  $m(A^0)$  less than the MeV range.
- <sup>2</sup> ALTMANN 95 looked for  $A^0$  decaying into  $e^+e^-$  from the Bugey 5 nuclear reactor. They obtain an upper limit on the  $A^0$  production rate of  $\omega(A^0)/\omega(\gamma) \times \text{B}(A^0 \rightarrow e^+e^-) < 10^{-16}$  for  $m_{A^0} = 1.5$  MeV at 90% CL. The limit is weaker for heavier  $A^0$ . In the case of a standard axion, this limit excludes a mass in the range  $2m_e < m_{A^0} < 4.8$  MeV at 90% CL. See Fig. 5 of their paper for exclusion limits of axion-like resonances  $Z^0$  in the  $(m_{X^0}, f_{X^0})$  plane.
- <sup>3</sup> KETOV 86 searched for  $A^0$  at the Rovno nuclear power plant. They found an upper limit on the  $A^0$  production probability of  $0.8 [100 \text{ keV}/m_{A^0}]^6 \times 10^{-6}$  per fission. In the standard axion model, this corresponds to  $m_{A^0} > 150$  keV. Not valid for  $m_{A^0} \gtrsim 1$  MeV.
- <sup>4</sup> KOCH 86 searched for  $A^0 \rightarrow \gamma\gamma$  at nuclear power reactor Biblis A. They found an upper limit on the  $A^0$  production rate of  $\omega(A^0)/\omega(\gamma(M1)) < 1.5 \times 10^{-10}$  (CL=95%). Standard axion with  $m_{A^0} = 250$  keV gives  $10^{-5}$  for the ratio. Not valid for  $m_{A^0} > 1022$  keV.
- <sup>5</sup> DATAR 82 looked for  $A^0 \rightarrow 2\gamma$  in neutron capture ( $n p \rightarrow d A^0$ ) at Tarapur 500 MW reactor. Sensitive to sum of  $l = 0$  and  $l = 1$  amplitudes. With ZEHNDER 81 [ $(l = 0) - (l = 1)$ ] result, assert nonexistence of standard  $A^0$ .
- <sup>6</sup> VUILLEUMIER 81 is at Grenoble reactor. Set limit  $m_{A^0} < 280$  keV.

### $A^0$ (Axion) and Other Light Boson ( $X^0$ ) Searches in Nuclear Transitions

Limits are for branching ratio.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 8.5 \times 10^{-6}$	90	1 DERBIN 02	CNTR	$^{125m}\text{Te}$ decay
		2 DEBOER 97c	RVUE	M1 transitions
$< 5.5 \times 10^{-10}$	95	3 TSUNODA 95	CNTR	$^{252}\text{Cf}$ fission, $A^0 \rightarrow e e$
$< 1.2 \times 10^{-6}$	95	4 MINOWA 93	CNTR	$^{139}\text{La}^* \rightarrow ^{139}\text{La} A^0$
$< 2 \times 10^{-4}$	90	5 HICKS 92	CNTR	$^{35}\text{S}$ decay, $A^0 \rightarrow \gamma\gamma$
$< 1.5 \times 10^{-9}$	95	6 ASANUMA 90	CNTR	$^{241}\text{Am}$ decay
$< (0.4-10) \times 10^{-3}$	95	7 DEBOER 90	CNTR	$^8\text{Be}^* \rightarrow ^8\text{Be} A^0$ , $A^0 \rightarrow e^+e^-$ , $^{16}\text{O}^* \rightarrow ^{16}\text{O} X^0$ , $X^0 \rightarrow e^+e^-$
$< (0.2-1) \times 10^{-3}$	90	8 BINI 89	CNTR	$^{16}\text{O}^* \rightarrow ^{16}\text{O} X^0$ , $X^0 \rightarrow e^+e^-$
		9 AVIGNONE 88	CNTR	$\text{Cu}^* \rightarrow \text{Cu} A^0 (A^0 \rightarrow 2\gamma, A^0 e \rightarrow \gamma e, A^0 Z \rightarrow \gamma Z)$
$< 1.5 \times 10^{-4}$	90	10 DATAR 88	CNTR	$^{12}\text{C}^* \rightarrow ^{12}\text{C} A^0$ , $A^0 \rightarrow e^+e^-$
$< 5 \times 10^{-3}$	90	11 DEBOER 88c	CNTR	$^{16}\text{O}^* \rightarrow ^{16}\text{O} X^0$ , $X^0 \rightarrow e^+e^-$
$< 3.4 \times 10^{-5}$	95	12 DOEHNER 88	SPEC	$2\text{H}^*, A^0 \rightarrow e^+e^-$
$< 4 \times 10^{-4}$	95	13 SAVAGE 88	CNTR	Nuclear decay (isovector)
$< 3 \times 10^{-3}$	95	13 SAVAGE 88	CNTR	Nuclear decay (isoscalar)
$< 10.6 \times 10^{-2}$	90	14 HALLIN 86	SPEC	$^6\text{Li}$ isovector decay
$< 10.8$	90	14 HALLIN 86	SPEC	$^{10}\text{B}$ isoscalar decays
$< 2.2$	90	14 HALLIN 86	SPEC	$^{14}\text{N}$ isoscalar decays
$< 4 \times 10^{-4}$	90	15 SAVAGE 86b	CNTR	$^{14}\text{N}^*$
		16 ANANEV 85	CNTR	$\text{Li}^*, \text{deut}^* A^0 \rightarrow 2\gamma$
		17 CAVAGNAC 83	CNTR	$^{97}\text{Nb}^*, \text{deut}^* \text{ transition } A^0 \rightarrow 2\gamma$
		18 ALEKSEEV 82b	CNTR	$\text{Li}^*, \text{deut}^* \text{ transition } A^0 \rightarrow 2\gamma$
		19 LEHMANN 82	CNTR	$\text{Cu}^* \rightarrow \text{Cu} A^0 (A^0 \rightarrow 2\gamma)$
		20 ZEHNDER 82	CNTR	$\text{Li}^*, \text{Nb}^* \text{ decay, } n\text{-capt.}$
		21 ZEHNDER 81	CNTR	$\text{Ba}^* \rightarrow \text{Ba} A^0 (A^0 \rightarrow 2\gamma)$
		22 CALAPRICE 79		Carbon

See key on page 601

## Gauge & Higgs Boson Particle Listings

### Axions ( $A^0$ ) and Other Very Light Bosons

- <sup>1</sup> DERBIN 02 looked for the axion emission in an M1 transition in  $^{125m}\text{Te}$  decay. They looked for a possible presence of a shifted energy spectrum in gamma rays due to the undetected axion.
- <sup>2</sup> DEBOER 97C reanalyzed the existent data on Nuclear M1 transitions and find that a 9 MeV boson decaying into  $e^+e^-$  would explain the excess of events with large opening angles. See also DEBOER 01 for follow-up experiments.
- <sup>3</sup> TSUNODA 95 looked for axion emission when  $^{252}\text{Cf}$  undergoes a spontaneous fission, with the axion decaying into  $e^+e^-$ . The bound is for  $m_{A^0}=40$  MeV. It improves to  $2.5 \times 10^{-5}$  for  $m_{A^0}=200$  MeV.
- <sup>4</sup> MINOWA 93 studied chain process,  $^{139}\text{Ce} \rightarrow ^{139}\text{La}^*$  by electron capture and M1 transition of  $^{139}\text{La}^*$  to the ground state. It does not assume decay modes of  $A^0$ . The bound applies for  $m_{A^0} < 166$  keV.
- <sup>5</sup> HICKS 92 bound is applicable for  $\tau_{X^0} < 4 \times 10^{-11}$  sec.
- <sup>6</sup> The ASANUMA 90 limit is for the branching fraction of  $X^0$  emission per  $^{241}\text{Am}$   $\alpha$  decay and valid for  $\tau_{X^0} < 3 \times 10^{-11}$  s.
- <sup>7</sup> The DEBOER 90 limit is for the branching ratio  $^8\text{Be}^* (18.15 \text{ MeV}, 1^+) \rightarrow ^8\text{Be} A^0, A^0 \rightarrow e^+e^-$  for the mass range  $m_{A^0} = 4\text{--}15$  MeV.
- <sup>8</sup> The BINI 89 limit is for the branching fraction of  $^{16}\text{O}^* (6.05 \text{ MeV}, 0^+) \rightarrow ^{16}\text{O} X^0, X^0 \rightarrow e^+e^-$  for  $m_X = 1.5\text{--}3.1$  MeV.  $\tau_{X^0} \lesssim 10^{-11}$  s is assumed. The spin-parity of  $X$  is restricted to  $0^+$  or  $1^-$ .
- <sup>9</sup> AVIGNONE 88 looked for the 1115 keV transition  $C^* \rightarrow \text{Cu} A^0$ , either from  $A^0 \rightarrow 2\gamma$  in-flight decay or from the secondary  $A^0$  interactions by Compton and by Primakoff processes. Limits for axion parameters are obtained for  $m_{A^0} < 1.1$  MeV.
- <sup>10</sup> DATAR 88 rule out light pseudoscalar particle emission through its decay  $A^0 \rightarrow e^+e^-$  in the mass range 1.02–2.5 MeV and lifetime range  $10^{-13}\text{--}10^{-8}$  s. The above limit is for  $\tau = 5 \times 10^{-13}$  s and  $m = 1.7$  MeV; see the paper for the  $\tau$ - $m$  dependence of the limit.
- <sup>11</sup> The limit is for the branching fraction of  $^{16}\text{O}^* (6.05 \text{ MeV}, 0^+) \rightarrow ^{16}\text{O} X^0, X^0 \rightarrow e^+e^-$  against internal pair conversion for  $m_{X^0} = 1.7$  MeV and  $\tau_{X^0} < 10^{-11}$  s. Similar limits are obtained for  $m_{X^0} = 1.3\text{--}3.2$  MeV. The spin parity of  $X^0$  must be either  $0^+$  or  $1^-$ . The limit at 1.7 MeV is translated into a limit for the  $X^0$ -nucleon coupling constant:  $g_{X^0 NN}^2/4\pi < 2.3 \times 10^{-9}$ .
- <sup>12</sup> The DOEHNER 88 limit is for  $m_{A^0} = 1.7$  MeV,  $\tau(A^0) < 10^{-10}$  s. Limits less than  $10^{-4}$  are obtained for  $m_{A^0} = 1.2\text{--}2.2$  MeV.
- <sup>13</sup> SAVAGE 88 looked for  $A^0$  that decays into  $e^+e^-$  in the decay of the 9.17 MeV  $J^P = 2^+$  state in  $^{14}\text{N}$ , 17.64 MeV state  $J^P = 1^+$  in  $^8\text{Be}$ , and the 18.15 MeV state  $J^P = 1^+$  in  $^8\text{Be}$ . This experiment constrains the isovector coupling of  $A^0$  to hadrons, if  $m_{A^0} = (1.1 \rightarrow 2.6)$  MeV and the isoscalar coupling of  $A^0$  to hadrons, if  $m_{A^0} = (1.1 \rightarrow 2.6)$  MeV. Both limits are valid only if  $\tau(A^0) \lesssim 1 \times 10^{-11}$  s.
- <sup>14</sup> Limits are for  $\Gamma(A^0(1.8 \text{ MeV}))/\Gamma(\pi\text{M1})$ ; i.e., for 1.8 MeV axion emission normalized to the rate for internal emission of  $e^+e^-$  pairs. Valid for  $\tau_{A^0} < 2 \times 10^{-11}$  s.  $^6\text{Li}$  isovector decay data strongly disfavor PECCEI 86 model I, whereas the  $^{10}\text{B}$  and  $^{14}\text{N}$  isoscalar decay data strongly reject PECCEI 86 model II and III.
- <sup>15</sup> SAVAGE 86b looked for  $A^0$  that decays into  $e^+e^-$  in the decay of the 9.17 MeV  $J^P = 2^+$  state in  $^{14}\text{N}$ . Limit on the branching fraction is valid if  $\tau_{A^0} \lesssim 1 \times 10^{-11}$  s for  $m_{A^0} = (1.1\text{--}1.7)$  MeV. This experiment constrains the iso-vector coupling of  $A^0$  to hadrons.
- <sup>16</sup> ANANEV 85 with IBR-2 pulsed reactor exclude standard  $A^0$  at CL = 95% masses below 470 keV ( $\text{Li}^*$  decay) and below  $2m_e$  for deuteron\* decay.
- <sup>17</sup> CAVAGNAC 83 at Bugey reactor exclude axion at any  $m_{97}\text{Nb}^*$  decay and axion with  $m_{A^0}$  between 275 and 288 keV (deuteron\* decay).
- <sup>18</sup> ALEKSEEV 82 with IBR-2 pulsed reactor exclude standard  $A^0$  at CL = 95% mass-ranges  $m_{A^0} < 400$  keV ( $\text{Li}^*$  decay) and  $330 \text{ keV} < m_{A^0} < 2.2$  MeV. (deuteron\* decay).
- <sup>19</sup> LEHMANN 82 obtained  $A^0 \rightarrow 2\gamma$  rate  $< 6.2 \times 10^{-5}/\text{s}$  (CL = 95%) excluding  $m_{A^0}$  between 100 and 1000 keV.
- <sup>20</sup> ZEHNDER 82 used Gosgen 2.8GW light-water reactor to check  $A^0$  production. No  $2\gamma$  peak in  $\text{Li}^*$ ,  $\text{Nb}^*$  decay (both single  $p$  transition) nor in  $n$  capture (combined with previous  $\text{Ba}^*$  negative result) rules out standard  $A^0$ . Set limit  $m_{A^0} < 60$  keV for any  $A^0$ .
- <sup>21</sup> ZEHNDER 81 looked for  $\text{Ba}^* \rightarrow A^0 \text{Ba}$  transition with  $A^0 \rightarrow 2\gamma$ . Obtained  $2\gamma$  coincidence rate  $< 2.2 \times 10^{-5}/\text{s}$  (CL = 95%) excluding  $m_{A^0} > 160$  keV (or 200 keV depending on Higgs mixing). However, see BARROSO 81.
- <sup>22</sup> CALAPRICE 79 saw no axion emission from excited states of carbon. Sensitive to axion mass between 1 and 15 MeV.

#### $A^0$ (Axion) Limits from Its Electron Coupling

Limits are for  $\tau(A^0 \rightarrow e^+e^-)$ .

VALUE (s)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none $4 \times 10^{-16}\text{--}4.5 \times 10^{-12}$	90	<sup>1</sup> BROSS	91	BDMP $eN \rightarrow eA^0N$ ( $A^0 \rightarrow ee$ )
		<sup>2</sup> GUO	90	BDMP $eN \rightarrow eA^0N$ ( $A^0 \rightarrow ee$ )
		<sup>3</sup> BJORKEN	88	CALO $A \rightarrow e^+e^-$ or $2\gamma$
		<sup>4</sup> BLINOV	88	MD1 $ee \rightarrow eeA^0$ ( $A^0 \rightarrow ee$ )
none $1 \times 10^{-14}\text{--}1 \times 10^{-10}$	90	<sup>5</sup> RIORDAN	87	BDMP $eN \rightarrow eA^0N$ ( $A^0 \rightarrow ee$ )
none $1 \times 10^{-14}\text{--}1 \times 10^{-11}$	90	<sup>6</sup> BROWN	86	BDMP $eN \rightarrow eA^0N$ ( $A^0 \rightarrow ee$ )
none $6 \times 10^{-14}\text{--}9 \times 10^{-11}$	95	<sup>7</sup> DAVIER	86	BDMP $eN \rightarrow eA^0N$ ( $A^0 \rightarrow ee$ )
none $3 \times 10^{-13}\text{--}1 \times 10^{-7}$	90	<sup>8</sup> KONAKA	86	BDMP $eN \rightarrow eA^0N$ ( $A^0 \rightarrow ee$ )

- <sup>1</sup> The listed BROSS 91 limit is for  $m_{A^0} = 1.14$  MeV.  $B(A^0 \rightarrow e^+e^-) = 1$  assumed. Excluded domain in the  $\tau_{A^0}$ - $m_{A^0}$  plane extends up to  $m_{A^0} \approx 7$  MeV (see Fig. 5). Combining with electron  $g$ -2 constraint, axions coupling only to  $e^+e^-$  ruled out for  $m_{A^0} < 4.8$  MeV (90% CL).
- <sup>2</sup> GUO 90 use the same apparatus as BROWN 86 and improve the previous limit in the shorter lifetime region. Combined with  $g$ -2 constraint, axions coupling only to  $e^+e^-$  are ruled out for  $m_{A^0} < 2.7$  MeV (90% CL).
- <sup>3</sup> BJORKEN 88 reports limits on axion parameters ( $f_A, m_A, \tau_A$ ) for  $m_{A^0} < 200$  MeV from electron beam-dump experiment with production via Primakoff photoproduction, bremsstrahlung from electrons, and resonant annihilation of positrons on atomic electrons.
- <sup>4</sup> BLINOV 88 assume zero spin,  $m = 1.8$  MeV and lifetime  $< 5 \times 10^{-12}$  s and find  $\Gamma(A^0 \rightarrow \gamma\gamma)B(A^0 \rightarrow e^+e^-) < 2$  eV (CL=90%).
- <sup>5</sup> Assumes  $A^0\gamma\gamma$  coupling is small and hence Primakoff production is small. Their figure 2 shows limits on axions for  $m_{A^0} < 15$  MeV.
- <sup>6</sup> Uses electrons in hadronic showers from an incident 800 GeV proton beam. Limits for  $m_{A^0} < 15$  MeV are shown in their figure 3.
- <sup>7</sup>  $m_{A^0} = 1.8$  MeV assumed. The excluded domain in the  $\tau_{A^0}$ - $m_{A^0}$  plane extends up to  $m_{A^0} \approx 14$  MeV, see their figure 4.
- <sup>8</sup> The limits are obtained from their figure 3. Also given is the limit on the  $A^0\gamma\gamma$ - $A^0e^+e^-$  coupling plane by assuming Primakoff production.

#### Search for $A^0$ (Axion) Resonance in Bhabha Scattering

The limit is for  $\Gamma(A^0)[B(A^0 \rightarrow e^+e^-)]^2$ .

VALUE ( $10^{-3}$ eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.3$	97	<sup>1</sup> HALLIN	92	CNTR $m_{A^0} = 1.75\text{--}1.88$ MeV
none 0.0016–0.47	90	<sup>2</sup> HENDERSON	92c	CNTR $m_{A^0} = 1.5\text{--}1.86$ MeV
$< 2.0$	90	<sup>3</sup> WU	92	CNTR $m_{A^0} = 1.56\text{--}1.86$ MeV
$< 0.013$	95	TSERTOS	91	CNTR $m_{A^0} = 1.832$ MeV
none 0.19–3.3	95	<sup>4</sup> WIDMANN	91	CNTR $m_{A^0} = 1.78\text{--}1.92$ MeV
$< 5$	97	BAUER	90	CNTR $m_{A^0} = 1.832$ MeV
none 0.09–1.5	95	<sup>5</sup> JUDGE	90	CNTR $m_{A^0} = 1.832$ MeV, elastic
$< 1.9$	97	<sup>6</sup> TSERTOS	89	CNTR $m_{A^0} = 1.82$ MeV
$< (10\text{--}40)$	97	<sup>6</sup> TSERTOS	89	CNTR $m_{A^0} = 1.51\text{--}1.65$ MeV
$< (1\text{--}2.5)$	97	<sup>6</sup> TSERTOS	89	CNTR $m_{A^0} = 1.80\text{--}1.86$ MeV
$< 31$	95	LORENZ	88	CNTR $m_{A^0} = 1.646$ MeV
$< 94$	95	LORENZ	88	CNTR $m_{A^0} = 1.726$ MeV
$< 23$	95	LORENZ	88	CNTR $m_{A^0} = 1.782$ MeV
$< 19$	95	LORENZ	88	CNTR $m_{A^0} = 1.837$ MeV
$< 3.8$	97	<sup>7</sup> TSERTOS	88	CNTR $m_{A^0} = 1.832$ MeV
		<sup>8</sup> VANKLINKEN	88	CNTR
		<sup>9</sup> MAIER	87	CNTR
$< 2500$	90	<sup>10</sup> MILLS	87	CNTR $m_{A^0} = 1.8$ MeV
		<sup>10</sup> VONWIMMER.87	CNTR	

- <sup>1</sup> HALLIN 92 quote limits on lifetime,  $8 \times 10^{-14} \text{--} 5 \times 10^{-13}$  sec depending on mass, assuming  $B(A^0 \rightarrow e^+e^-) = 100\%$ . They say that TSERTOS 91 overstated their sensitivity by a factor of 3.
- <sup>2</sup> HENDERSON 92c exclude axion with lifetime  $\tau_{A^0} = 1.4 \times 10^{-12} \text{--} 4.0 \times 10^{-10}$  s, assuming  $B(A^0 \rightarrow e^+e^-) = 100\%$ . HENDERSON 92c also exclude a vector boson with  $\tau = 1.4 \times 10^{-12} \text{--} 6.0 \times 10^{-10}$  s.
- <sup>3</sup> WU 92 quote limits on lifetime  $> 3.3 \times 10^{-13}$  s assuming  $B(A^0 \rightarrow e^+e^-) = 100\%$ . They say that TSERTOS 89 overestimate the limit by a factor of  $\pi/2$ . WU 92 also quote a bound for vector boson,  $\tau > 8.2 \times 10^{-13}$  s.
- <sup>4</sup> WIDMANN 91 bound applies exclusively to the case  $B(A^0 \rightarrow e^+e^-) = 1$ , since the detection efficiency varies substantially as  $\Gamma(A^0)_{\text{total}}$  changes. See their Fig. 6.
- <sup>5</sup> JUDGE 90 excludes an elastic pseudoscalar  $e^+e^-$  resonance for  $4.5 \times 10^{-13} \text{ s} < \tau(A^0) < 7.5 \times 10^{-12} \text{ s}$  (95% CL) at  $m_{A^0} = 1.832$  MeV. Comparable limits can be set for  $m_{A^0} = 1.776\text{--}1.856$  MeV.
- <sup>6</sup> See also TSERTOS 88b in references.
- <sup>7</sup> The upper limit listed in TSERTOS 88 is too large by a factor of 4. See TSERTOS 88b, footnote 3.
- <sup>8</sup> VANKLINKEN 88 looked for relatively long-lived resonance ( $\tau = 10^{-10}\text{--}10^{-12}$  s). The sensitivity is not sufficient to exclude such a narrow resonance.
- <sup>9</sup> MAIER 87 obtained limits  $R\Gamma \lesssim 60$  eV (100 eV) at  $m_{A^0} \approx 1.64$  MeV (1.83 MeV) for energy resolution  $\Delta E_{\text{cm}} \approx 3$  keV, where  $R$  is the resonance cross section normalized to that of Bhabha scattering, and  $\Gamma = \Gamma_{e^+e^-}^2/\Gamma_{\text{total}}$ . For a discussion implying that  $\Delta E_{\text{cm}} \approx 10$  keV, see TSERTOS 89.
- <sup>10</sup> VONWIMMERSPERG 87 measured Bhabha scattering for  $E_{\text{cm}} = 1.37\text{--}1.86$  MeV and found a possible peak at 1.73 with  $\int \sigma dE_{\text{cm}} = 14.5 \pm 6.8$  keV-b. For a comment and a reply, see VANKLINKEN 88b and VONWIMMERSPERG 88. Also see CONNELL 88.

#### Search for $A^0$ (Axion) Resonance in $e^+e^- \rightarrow \gamma\gamma$

The limit is for  $\Gamma(A^0 \rightarrow e^+e^-)\Gamma(A^0 \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$

VALUE ( $10^{-3}$ eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

< 0.18	95	VO	94	CNTR	$m_{A^0}=1.1$ MeV
< 1.5	95	VO	94	CNTR	$m_{A^0}=1.4$ MeV
<12	95	VO	94	CNTR	$m_{A^0}=1.7$ MeV
< 6.6	95	<sup>1</sup> TRZASKA	91	CNTR	$m_{A^0}=1.8$ MeV
< 4.4	95	WIDMANN	91	CNTR	$m_{A^0}=1.78\text{--}1.92$ MeV
		<sup>2</sup> FOX	89	CNTR	
< 0.11	95	<sup>3</sup> MINOWA	89	CNTR	$m_{A^0}=1.062$ MeV
<33	97	CONNELL	88	CNTR	$m_{A^0}=1.580$ MeV
<42	97	CONNELL	88	CNTR	$m_{A^0}=1.642$ MeV
<73	97	CONNELL	88	CNTR	$m_{A^0}=1.782$ MeV
<79	97	CONNELL	88	CNTR	$m_{A^0}=1.832$ MeV

- <sup>1</sup> TRZASKA 91 also give limits in the range  $(6.6\text{--}30) \times 10^{-3}$  eV (95%CL) for  $m_{A^0} = 1.6\text{--}2.0$  MeV.
- <sup>2</sup> FOX 89 measured positron annihilation with an electron in the source material into two photons and found no signal at 1.062 MeV ( $< 9 \times 10^{-5}$  of two-photon annihilation at rest).
- <sup>3</sup> Similar limits are obtained for  $m_{A^0} = 1.045\text{--}1.085$  MeV.

### Search for $X^0$ (Light Boson) Resonance in $e^+e^- \rightarrow \gamma\gamma\gamma$

The limit is for  $\Gamma(X^0 \rightarrow e^+e^-)\Gamma(X^0 \rightarrow \gamma\gamma\gamma)/\Gamma_{\text{total}}$ .  $C$  invariance forbids spin-0  $X^0$  coupling to both  $e^+e^-$  and  $\gamma\gamma\gamma$ .

VALUE ( $10^{-3}$ eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
< 0.2	95	<sup>1</sup> VO	94	CNTR $m_{X^0}=1.1\text{--}1.9$ MeV
< 1.0	95	<sup>2</sup> VO	94	CNTR $m_{X^0}=1.1$ MeV
< 2.5	95	<sup>2</sup> VO	94	CNTR $m_{X^0}=1.4$ MeV
<120	95	<sup>2</sup> VO	94	CNTR $m_{X^0}=1.7$ MeV
< 3.8	95	<sup>3</sup> SKALSEY	92	CNTR $m_{X^0}=1.5$ MeV

- <sup>1</sup> VO 94 looked for  $X^0 \rightarrow \gamma\gamma\gamma$  decaying at rest. The precise limits depend on  $m_{X^0}$ . See Fig. 2(b) in paper.
- <sup>2</sup> VO 94 looked for  $X^0 \rightarrow \gamma\gamma\gamma$  decaying in flight.
- <sup>3</sup> SKALSEY 92 also give limits 4.3 for  $m_{X^0} = 1.54$  and 7.5 for 1.64 MeV. The spin of  $X^0$  is assumed to be one.

### Light Boson ( $X^0$ ) Search in Nonresonant $e^+e^-$ Annihilation at Rest

Limits are for the ratio of  $n\gamma + X^0$  production relative to  $\gamma\gamma$ .

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
< 4.2	90	<sup>1</sup> MITSUI	96	CNTR $\gamma X^0$
< 4	68	<sup>2</sup> SKALSEY	95	CNTR $\gamma X^0$
<40	68	<sup>3</sup> SKALSEY	95	RVUE $\gamma X^0$
< 0.18	90	<sup>4</sup> ADACHI	94	CNTR $\gamma\gamma X^0, X^0 \rightarrow \gamma\gamma$
< 0.26	90	<sup>5</sup> ADACHI	94	CNTR $\gamma\gamma X^0, X^0 \rightarrow \gamma\gamma$
< 0.33	90	<sup>6</sup> ADACHI	94	CNTR $\gamma X^0, X^0 \rightarrow \gamma\gamma\gamma$

- <sup>1</sup> MITSUI 96 looked for a monochromatic  $\gamma$ . The bound applies for a vector  $X^0$  with  $C=-1$  and  $m_{X^0} < 200$  keV. They derive an upper bound on  $eeX^0$  coupling and hence on the branching ratio  $B(o\text{-Ps} \rightarrow \gamma\gamma X^0) < 6.2 \times 10^{-6}$ . The bounds weaken for heavier  $X^0$ .
- <sup>2</sup> SKALSEY 95 looked for a monochromatic  $\gamma$  without an accompanying  $\gamma$  in  $e^+e^-$  annihilation. The bound applies for scalar and vector  $X^0$  with  $C=-1$  and  $m_{X^0} = 100\text{--}1000$  keV.
- <sup>3</sup> SKALSEY 95 reinterpreted the bound on  $\gamma A^0$  decay of  $o\text{-Ps}$  by ASAI 91 where 3% of delayed annihilations are not from  $^3S_1$  states. The bound applies for scalar and vector  $X^0$  with  $C=-1$  and  $m_{X^0} = 0\text{--}800$  keV.
- <sup>4</sup> ADACHI 94 looked for a peak in the  $\gamma\gamma$  invariant mass distribution in  $\gamma\gamma\gamma\gamma$  production from  $e^+e^-$  annihilation. The bound applies for  $m_{X^0} = 70\text{--}800$  keV.
- <sup>5</sup> ADACHI 94 looked for a peak in the missing-mass mass distribution in  $\gamma\gamma$  channel, using  $\gamma\gamma\gamma\gamma$  production from  $e^+e^-$  annihilation. The bound applies for  $m_{X^0} < 800$  keV.
- <sup>6</sup> ADACHI 94 looked for a peak in the missing mass distribution in  $\gamma\gamma\gamma$  channel, using  $\gamma\gamma\gamma\gamma$  production from  $e^+e^-$  annihilation. The bound applies for  $m_{X^0} = 200\text{--}900$  keV.

### Searches for Goldstone Bosons ( $X^0$ )

(Including Horizontal Bosons and Majorons.) Limits are for branching ratios.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
<9 $\times 10^{-6}$	90	<sup>1</sup> BAYES	15	TWST $\mu^+ \rightarrow e^+X^0$ , Familon
		<sup>2</sup> LATTANZI	13	COSM Majoron dark matter decay
		<sup>3</sup> LESSA	07	RVUE Meson, $\ell$ decays to Majoron
		<sup>4</sup> DIAZ	98	THEO $H^0 \rightarrow X^0X^0, A^0 \rightarrow X^0X^0X^0$ , Majoron
		<sup>5</sup> BOBRAKOV	91	Electron quasi-magnetic interaction
<3.3 $\times 10^{-2}$	95	<sup>6</sup> ALBRECHT	90E	ARG $\tau \rightarrow \mu X^0$ , Familon
<1.8 $\times 10^{-2}$	95	<sup>6</sup> ALBRECHT	90E	ARG $\tau \rightarrow e X^0$ , Familon
<6.4 $\times 10^{-9}$	90	<sup>7</sup> ATIYA	90	B787 $K^+ \rightarrow \pi^+X^0$ , Familon
<1.1 $\times 10^{-9}$	90	<sup>8</sup> BOLTON	88	CBOX $\mu^+ \rightarrow e^+\gamma X^0$ , Familon

		<sup>9</sup> CHANDA	88	ASTR Sun, Majoron
		<sup>10</sup> CHOI	88	ASTR Majoron, SN 1987A
<5 $\times 10^{-6}$	90	<sup>11</sup> PICCIOTTO	88	CNTR $\pi \rightarrow e\nu X^0$ , Majoron
<1.3 $\times 10^{-9}$	90	<sup>12</sup> GOLDMAN	87	CNTR $\mu \rightarrow e\gamma X^0$ , Familon
<3 $\times 10^{-4}$	90	<sup>13</sup> BRYMAN	86B	RVUE $\mu \rightarrow eX^0$ , Familon
<1 $\times 10^{-10}$	90	<sup>14</sup> EICHLER	86	SPEC $\mu^+ \rightarrow e^+X^0$ , Familon
<2.6 $\times 10^{-6}$	90	<sup>15</sup> JODIDIO	86	SPEC $\mu^+ \rightarrow e^+X^0$ , Familon
		<sup>16</sup> BALTRUSAITIS	85	MRK3 $\tau \rightarrow \ell X^0$ , Familon
		<sup>17</sup> DICUS	83	COSM $\nu(\text{hvy}) \rightarrow \nu(\text{light}) X^0$

- <sup>1</sup> BAYES 15 limits are the average over  $m_{X^0} = 13\text{--}80$  MeV for the isotropic decay distribution of positrons. See their Fig. 4 and Table II for the mass-dependent limits as well as the dependence on the decay anisotropy. In particular, they find a limit  $< 58 \times 10^{-6}$  at 90% CL for massless familons and for the same asymmetry as normal moon decay, a case not covered by JODIDIO 86.
- <sup>2</sup> LATTANZI 13 use WMAP 9 year data as well as X-ray and  $\gamma$ -ray observations to derive limits on decaying majoron dark matter. A limit on the decay width  $\Gamma(X^0 \rightarrow \nu\bar{\nu}) < 6.4 \times 10^{-19} \text{ s}^{-1}$  at 95% CL is found if majorons make up all of the dark matter.
- <sup>3</sup> LESSA 07 consider decays of the form Meson  $\rightarrow \ell\nu\text{Majoron}$  and  $\ell \rightarrow \ell'\nu\bar{\nu}\text{Majoron}$  and use existing data to derive limits on the neutrino-Majoron Yukawa couplings  $g_{\alpha\beta}$  ( $\alpha, \beta = e, \mu, \tau$ ). Their best limits are  $|g_{e\alpha}|^2 < 5.5 \times 10^{-6}$ ,  $|g_{\mu\alpha}|^2 < 4.5 \times 10^{-5}$ ,  $|g_{\tau\alpha}|^2 < 5.5 \times 10^{-2}$  at CL = 90%.
- <sup>4</sup> DIAZ 98 studied models of spontaneously broken lepton number with both singlet and triplet Higgses. They obtain limits on the parameter space from invisible decay  $Z \rightarrow H^0 A^0 \rightarrow X^0 X^0 X^0 X^0$  and  $e^+e^- \rightarrow Z H^0$  with  $H^0 \rightarrow X^0 X^0$ .
- <sup>5</sup> BOBRAKOV 91 searched for anomalous magnetic interactions between polarized electrons expected from the exchange of a massless pseudoscalar boson (arion). A limit  $x_e^2 < 2 \times 10^{-4}$  (95%CL) is found for the effective anomalous magneton parametrized as  $x_e(G_F/8\pi\sqrt{2})^{1/2}$ .
- <sup>6</sup> ALBRECHT 90E limits are for  $B(\tau \rightarrow \ell X^0)/B(\tau \rightarrow \ell\nu\bar{\nu})$ . Valid for  $m_{X^0} < 100$  MeV. The limits rise to 7.1% (for  $\mu$ ), 5.0% (for  $e$ ) for  $m_{X^0} = 500$  MeV.
- <sup>7</sup> ATIYA 90 limit is for  $m_{X^0} = 0$ . The limit  $B < 1 \times 10^{-8}$  holds for  $m_{X^0} < 95$  MeV.
- For the reduction of the limit due to finite lifetime of  $X^0$ , see their Fig. 3.
- <sup>8</sup> BOLTON 88 limit corresponds to  $F > 3.1 \times 10^9$  GeV, which does not depend on the chirality property of the coupling.
- <sup>9</sup> CHANDA 88 find  $v_T < 10$  MeV for the weak-triplet Higgs vacuum expectation value in Gelmini-Roncadelli model, and  $v_S > 5.8 \times 10^6$  GeV in the singlet Majoron model.
- <sup>10</sup> CHOI 88 used the observed neutrino flux from the supernova SN1987A to exclude the neutrino Majoron Yukawa coupling  $h$  in the range  $2 \times 10^{-5} < h < 3 \times 10^{-4}$  for the interaction  $L_{\text{int}} = \frac{1}{2}ih\bar{\psi}_\nu\gamma_5\psi_\nu\phi_X$ . For several families of neutrinos, the limit applies for  $(\Sigma h_i^4)^{1/4}$ .
- <sup>11</sup> PICCIOTTO 88 limit applies when  $m_{X^0} < 55$  MeV and  $\tau_{X^0} > 2\text{ns}$ , and it decreases to  $4 \times 10^{-7}$  at  $m_{X^0} = 125$  MeV, beyond which no limit is obtained.
- <sup>12</sup> GOLDMAN 87 limit corresponds to  $F > 2.9 \times 10^9$  GeV for the family symmetry breaking scale from the Lagrangian  $L_{\text{int}} = (1/F)\bar{\psi}_\mu\gamma^\mu(a + b\gamma_5)\psi_e\partial_\mu\phi_X$  with  $a^2 + b^2 = 1$ . This is not as sensitive as the limit  $F > 9.9 \times 10^9$  GeV derived from the search for  $\mu^+ \rightarrow e^+X^0$  by JODIDIO 86, but does not depend on the chirality property of the coupling.
- <sup>13</sup> Limits are for  $\Gamma(\mu \rightarrow eX^0)/\Gamma(\mu \rightarrow e\nu\bar{\nu})$ . Valid when  $m_{X^0} = 0\text{--}93.4, 98.1\text{--}103.5$  MeV.
- <sup>14</sup> EICHLER 86 looked for  $\mu^+ \rightarrow e^+X^0$  followed by  $X^0 \rightarrow e^+e^-$ . Limits on the branching fraction depend on the mass and lifetime of  $X^0$ . The quoted limits are valid when  $\tau_{X^0} \lesssim 3. \times 10^{-10} \text{ s}$  if the decays are kinematically allowed.
- <sup>15</sup> JODIDIO 86 corresponds to  $F > 9.9 \times 10^9$  GeV for the family symmetry breaking scale with the parity-conserving effective Lagrangian  $L_{\text{int}} = (1/F)\bar{\psi}_\mu\gamma^\mu\psi_e\partial^\mu\phi_X$ .
- <sup>16</sup> BALTRUSAITIS 85 search for light Goldstone boson( $X^0$ ) of broken  $U(1)$ . CL = 95% limits are  $B(\tau \rightarrow \mu^+X^0)/B(\tau \rightarrow \mu^+\nu\bar{\nu}) < 0.125$  and  $B(\tau \rightarrow e^+X^0)/B(\tau \rightarrow e^+\nu\bar{\nu}) < 0.04$ . Inferred limit for the symmetry breaking scale is  $m > 3000$  TeV.
- <sup>17</sup> The primordial heavy neutrino must decay into  $\nu$  and familon,  $f_A$ , early so that the red-shifted decay products are below critical density, see their table. In addition,  $K \rightarrow \pi f_A$  and  $\mu \rightarrow e f_A$  are unseen. Combining these excludes  $m_{\text{heavy}\nu}$  between  $5 \times 10^{-5}$  and  $5 \times 10^{-4}$  MeV ( $\mu$  decay) and  $m_{\text{heavy}\nu}$  between  $5 \times 10^{-5}$  and 0.1 MeV ( $K$ -decay).

### Majoron Searches in Neutrinoless Double $\beta$ Decay

Limits are for the half-life of neutrinoless  $\beta\beta$  decay with a Majoron emission.

No experiment currently claims any such evidence. Only the best or comparable limits for each isotope are reported. Also see the reviews ZUBER 98 and FAESSLER 98B.

$t_{1/2}(10^{21} \text{ yr})$	CL%	ISOTOPE	TRANSITION	METHOD	DOCUMENT ID
>7200		<b>90 128Te</b>		<b>CNTR</b>	<sup>1</sup> BERNATOW... 92
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
> 420	90	<sup>76</sup> Ge	$0\nu 1\chi$	GERDA	<sup>2</sup> AGOSTINI 15A
> 400	90	<sup>100</sup> Mo	$0\nu 1\chi$	NEMO-3	<sup>3</sup> ARNOLD 15
>1200	90	<sup>136</sup> Xe	$0\nu 1\chi$	EXO-200	<sup>4</sup> ALBERT 14A
>2600	90	<sup>136</sup> Xe	$0\nu 1\chi$	KamLAND-Zen	<sup>5</sup> GANDO 12
> 16	90	<sup>130</sup> Te	$0\nu 1\chi$	NEMO-3	<sup>6</sup> ARNOLD 11
> 1.9	90	<sup>96</sup> Zr	$2\nu 1\chi$	NEMO-3	<sup>7</sup> ARGYRIADES 10
> 1.52	90	<sup>150</sup> Nd	$0\nu 1\chi$	NEMO-3	<sup>8</sup> ARGYRIADES 09
> 27	90	<sup>100</sup> Mo	$0\nu 1\chi$	NEMO-3	<sup>9</sup> ARNOLD 06
> 15	90	<sup>82</sup> Se	$0\nu 1\chi$	NEMO-3	<sup>10</sup> ARNOLD 06
> 14	90	<sup>100</sup> Mo	$0\nu 1\chi$	NEMO-3	<sup>11</sup> ARNOLD 04
> 12	90	<sup>82</sup> Se	$0\nu 1\chi$	NEMO-3	<sup>12</sup> ARNOLD 04
> 2.2	90	<sup>130</sup> Te	$0\nu 1\chi$	Cryog. det.	<sup>13</sup> ARNABOLDI 03
> 0.9	90	<sup>130</sup> Te	$0\nu 2\chi$	Cryog. det.	<sup>14</sup> ARNABOLDI 03

See key on page 601

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

> 8	90	$^{116}\text{Cd}$	$0\nu 1\chi$	$\text{CdWO}_4$ scint.	15	DANEVICH	03
> 0.8	90	$^{116}\text{Cd}$	$0\nu 2\chi$	$\text{CdWO}_4$ scint.	16	DANEVICH	03
> 500	90	$^{136}\text{Xe}$	$0\nu 1\chi$	Liquid Xe Scint.	17	BERNABEI	02p
> 5.8	90	$^{100}\text{Mo}$	$0\nu 1\chi$	ELEGANT V	18	FUSHIMI	02
> 0.32	90	$^{100}\text{Mo}$	$0\nu 1\chi$	Liq. Ar ioniz.	19	ASHITKOV	01
> 0.0035	90	$^{160}\text{Gd}$	$0\nu 1\chi$	$^{160}\text{Gd}_2\text{SiO}_5:\text{Ce}$	20	DANEVICH	01
> 0.013	90	$^{160}\text{Gd}$	$0\nu 2\chi$	$^{160}\text{Gd}_2\text{SiO}_5:\text{Ce}$	21	DANEVICH	01
> 2.3	90	$^{82}\text{Se}$	$0\nu 1\chi$	NEMO 2	22	ARNOLD	00
> 0.31	90	$^{96}\text{Zr}$	$0\nu 1\chi$	NEMO 2	23	ARNOLD	00
> 0.63	90	$^{82}\text{Se}$	$0\nu 2\chi$	NEMO 2	24	ARNOLD	00
> 0.063	90	$^{96}\text{Zr}$	$0\nu 2\chi$	NEMO 2	24	ARNOLD	00
> 0.16	90	$^{100}\text{Mo}$	$0\nu 2\chi$	NEMO 2	24	ARNOLD	00
> 2.4	90	$^{82}\text{Se}$	$0\nu 1\chi$	NEMO 2	25	ARNOLD	98
> 7.2	90	$^{136}\text{Xe}$	$0\nu 2\chi$	TPC	26	LUESCHER	98
> 7.91	90	$^{76}\text{Ge}$		SPEC	27	GUENTHER	96
> 17	90	$^{76}\text{Ge}$		CNTR		BECK	93

- 1 BERNATOWICZ 92 studied double- $\beta$  decays of  $^{128}\text{Te}$  and  $^{130}\text{Te}$ , and found the ratio  $\tau(^{130}\text{Te})/\tau(^{128}\text{Te}) = (3.52 \pm 0.11) \times 10^{-4}$  in agreement with relatively stable theoretical predictions. The bound is based on the requirement that Majoron-emitting decay cannot be larger than the observed double-beta rate of  $^{128}\text{Te}$  of  $(7.7 \pm 0.4) \times 10^{24}$  year. We calculated 90% CL limit as  $(7.7-1.28 \times 0.4=7.2) \times 10^{24}$ .
- 2 AGOSTINI 15A analyze a 20.3 kg yr of data set of the GERDA calorimeter to determine  $g_{\nu\chi} < 3.4-8.7 \times 10^{-5}$  on the Majoron-neutrino coupling constant. The range reflects the spread of the nuclear matrix elements.
- 3 ARNOLD 15 use the NEMO-3 tracking calorimeter with 3.43 kg yr exposure to determine the limit on Majoron emission. The limit corresponds to  $g_{\nu\chi} < 1.6-3.0 \times 10^{-4}$ . The spread reflects different nuclear matrix elements. Supersedes ARNOLD 06.
- 4 ALBERT 14A utilize 100 kg yr of exposure of the EXO-200 tracking calorimeter to place a limit on the  $g_{\nu\chi} < 0.8-1.7 \times 10^{-5}$  on the Majoron-neutrino coupling constant. The range reflects the spread of the nuclear matrix elements.
- 5 GANDO 12 use the KamLAND-Zen detector to obtain the limit on the  $0\nu\chi$  decay with Majoron emission. It implies that the coupling constant  $g_{\nu\chi} < 0.8-1.6 \times 10^{-5}$  depending on the nuclear matrix elements used.
- 6 ARNOLD 11 use the NEMO-3 detector to obtain the reported limit on Majoron emission. It implies that the coupling constant  $g_{\nu\chi} < 0.6-1.6 \times 10^{-4}$  depending on the nuclear matrix element used. Supersedes ARNOLDI 03.
- 7 ARGYRIADES 10 use the NEMO-3 tracking detector and  $^{96}\text{Zr}$  to derive the reported limit. No limit for the Majoron electron coupling is given.
- 8 ARGYRIADES 09 use  $^{150}\text{Nd}$  data taken with the NEMO-3 tracking detector. The reported limit corresponds to  $\langle g_{\nu\chi} \rangle < 1.7-3.0 \times 10^{-4}$  using a range of nuclear matrix elements that include the effect of nuclear deformation.
- 9 ARNOLD 06 use  $^{100}\text{Mo}$  data taken with the NEMO-3 tracking detector. The reported limit corresponds to  $\langle g_{\nu\chi} \rangle < (0.4-1.8) \times 10^{-4}$  using a range of matrix element calculations. Superseded by ARNOLD 15.
- 10 NEMO-3 tracking calorimeter is used in ARNOLD 06. Reported half-life limit for  $^{82}\text{Se}$  corresponds to  $\langle g_{\nu\chi} \rangle < (0.66-1.9) \times 10^{-4}$  using a range of matrix element calculations. Supersedes ARNOLD 04.
- 11 ARNOLD 04 use the NEMO-3 tracking detector. The limit corresponds to  $\langle g_{\nu\chi} \rangle < (0.5-0.9)10^{-4}$  using the matrix elements of SIMKOVIĆ 99, STOICA 01 and CIVITARESE 03. Supersedes by ARNOLD 06.
- 12 ARNOLD 04 use the NEMO-3 tracking detector. The limit corresponds to  $\langle g_{\nu\chi} \rangle < (0.7-1.6)10^{-4}$  using the matrix elements of SIMKOVIĆ 99, STOICA 01 and CIVITARESE 03.
- 13 Supersedes ALESSANDRELLO 00. Array of  $\text{TeO}_2$  crystals in high resolution cryogenic calorimeter. Some enriched in  $^{130}\text{Te}$ . Derive  $\langle g_{\nu\chi} \rangle < 17-33 \times 10^{-5}$  depending on matrix element.
- 14 Supersedes ALESSANDRELLO 00. Cryogenic calorimeter search.
- 15 Limit for the  $0\nu\chi$  decay with Majoron emission of  $^{116}\text{Cd}$  using enriched  $\text{CdWO}_4$  scintillators.  $\langle g_{\nu\chi} \rangle < 4.6-8.1 \times 10^{-5}$  depending on the matrix element. Supersedes DANEVICH 00.
- 16 Limit for the  $0\nu 2\chi$  decay of  $^{116}\text{Cd}$ . Supersedes DANEVICH 00.
- 17 BERNABEI 02p obtain limit for  $0\nu\chi$  decay with Majoron emission of  $^{136}\text{Xe}$  using liquid Xe scintillation detector. They derive  $\langle g_{\nu\chi} \rangle < 2.0-3.0 \times 10^{-5}$  with several nuclear matrix elements.
- 18 Replaces TANAKA 93. FUSHIMI 02 derive half-life limit for the  $0\nu\chi$  decay by means of tracking calorimeter ELEGANT V. Considering various matrix element calculations, a range of limits for the Majoron-neutrino coupling is given:  $\langle g_{\nu\chi} \rangle < (6.3-360) \times 10^{-5}$ .
- 19 ASHITKOV 01 result for  $0\nu\chi$  of  $^{100}\text{Mo}$  is less stringent than ARNOLD 00.
- 20 DANEVICH 01 obtain limit for the  $0\nu\chi$  decay with Majoron emission of  $^{160}\text{Gd}$  using  $\text{Gd}_2\text{SiO}_5:\text{Ce}$  crystal scintillators.
- 21 DANEVICH 01 obtain limit for the  $0\nu 2\chi$  decay with 2 Majoron emission of  $^{160}\text{Gd}$ .
- 22 ARNOLD 00 reports limit for the  $0\nu\chi$  decay with Majoron emission derived from tracking calorimeter NEMO 2. Using  $^{82}\text{Se}$  source:  $\langle g_{\nu\chi} \rangle < 1.6 \times 10^{-4}$ . Matrix element from GUENTHER 96.
- 23 Using  $^{96}\text{Zr}$  source:  $\langle g_{\nu\chi} \rangle < 2.6 \times 10^{-4}$ . Matrix element from ARNOLD 99.
- 24 ARNOLD 00 reports limit for the  $0\nu 2\chi$  decay with two Majoron emission derived from tracking calorimeter NEMO 2.
- 25 ARNOLD 98 determine the limit for  $0\nu\chi$  decay with Majoron emission of  $^{82}\text{Se}$  using the NEMO-2 tracking detector. They derive  $\langle g_{\nu\chi} \rangle < 2.3-4.3 \times 10^{-4}$  with several nuclear matrix elements.
- 26 LUESCHER 98 report a limit for the  $0\nu$  decay with Majoron emission of  $^{136}\text{Xe}$  using Xe TPC. This result is more stringent than BARABASH 89. Using the matrix elements of ENGEL 88, they obtain a limit on  $\langle g_{\nu\chi} \rangle$  of  $2.0 \times 10^{-4}$ .
- 27 See Table I in GUENTHER 96 for limits on the Majoron coupling in different models.

### Invisible $A^0$ (Axion) MASS LIMITS from Astrophysics and Cosmology

$v_1 = v_2$  is usually assumed ( $v_i$  = vacuum expectation values). For a review of these limits, see RAFFELT 91 and TURNER 90. In the comment lines below, D and K refer to DFSZ and KSVZ axion types, discussed in the above minireview.

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.67	95	1 ARCHIDIACO...13A	COSM	K, hot dark matter
none $0.7-3 \times 10^5$		2 CADAMURO 11	COSM	D abundance
<105	90	3 DERBIN 11A	CNTR	D, solar axion
		4 ANDRIAMON...10	CAST	K, solar axions
< 0.72	95	5 HANNESTAD 10	COSM	K, hot dark matter
		6 ANDRIAMON...09	CAST	K, solar axions
<191	90	7 DERBIN 09A	CNTR	K, solar axions
<334	95	8 KEKEZ 09	HPGE	K, solar axions
< 1.02	95	9 HANNESTAD 08	COSM	K, hot dark matter
< 1.2	95	10 HANNESTAD 07	COSM	K, hot dark matter
< 0.42	95	11 MELCHIORRI 07A	COSM	K, hot dark matter
< 1.05	95	12 HANNESTAD 05A	COSM	K, hot dark matter
3 to 20		13 MOROI 98	COSM	K, hot dark matter
< 0.007		14 BORISOV 97	ASTR	D, neutron star
< 4		15 KACHELRIESS 97	ASTR	D, neutron star cooling
<(0.5-6) $\times 10^{-3}$		16 KEIL 97	ASTR	SN 1987A
< 0.018		17 RAFFELT 95	ASTR	D, red giant
< 0.010		18 ALTHERR 94	ASTR	D, red giants, white dwarfs
< 0.01		19 CHANG 93	ASTR	K, SN 1987A
< 0.03		WANG 92	ASTR	D, white dwarf
none 3-8		WANG 92C	ASTR	D, C-O burning
		20 BERSHADY 91	ASTR	D, K, intergalactic light
< 10		21 KIM 91C	COSM	D, K, mass density of the universe, super-symmetry
< 1 $\times 10^{-3}$		22 RAFFELT 91B	ASTR	D,K, SN 1987A
none $10^{-3}-3$		23 RESSELL 91	ASTR	K, intergalactic light
		BURROWS 90	ASTR	D,K, SN 1987A
< 0.02		24 ENGEL 90	ASTR	D,K, SN 1987A
< 1 $\times 10^{-3}$		25 RAFFELT 90B	ASTR	D, red giant
<(1.4-10) $\times 10^{-3}$		26 BURROWS 89	ASTR	D,K, SN 1987A
< 3.6 $\times 10^{-4}$		27 ERICSON 89	ASTR	D,K, SN 1987A
< 12		28 MAYLE 89	ASTR	D,K, SN 1987A
< 1 $\times 10^{-3}$		CHANDA 88	ASTR	D, Sun
		RAFFELT 88	ASTR	D,K, SN 1987A
< 0.07		29 RAFFELT 88B	ASTR	red giant
< 0.7		FRIEMAN 87	ASTR	D, red giant
< 2-5		30 RAFFELT 87	ASTR	K, red giant
< 0.01		TURNER 87	COSM	K, thermal production
< 0.06		31 DEARBORN 86	ASTR	D, red giant
< 0.7		RAFFELT 86	ASTR	D, red giant
< 0.03		32 RAFFELT 86	ASTR	K, red giant
< 1		RAFFELT 86B	ASTR	D, white dwarf
< 0.003-0.02		33 KAPLAN 85	ASTR	K, red giant
> 1 $\times 10^{-5}$		IWAMOTO 84	ASTR	D, K, neutron star
		ABBOTT 83	COSM	D,K, mass density of the universe
> 1 $\times 10^{-5}$		DINE 83	COSM	D,K, mass density of the universe
< 0.04		ELLIS 83B	ASTR	D, red giant
> 1 $\times 10^{-5}$		PRESKILL 83	COSM	D,K, mass density of the universe
< 0.1		BARROSO 82	ASTR	D, red giant
< 1		34 FUKUGITA 82	ASTR	D, stellar cooling
< 0.07		FUKUGITA 82B	ASTR	D, red giant

- 1 ARCHIDIACONO 13A is analogous to HANNESTAD 05A. The limit is based on the CMB temperature power spectrum of the Planck data, the CMB polarization from the WMAP 9-yr data, the matter power spectrum from SDSS-DR7, and the local Hubble parameter measurement by the Carnegie Hubble program.
- 2 CADAMURO 11 use the deuterium abundance to show that the  $m_{A^0}$  range 0.7 eV – 300 keV is excluded for axions, complementing HANNESTAD 10.
- 3 DERBIN 11A look for solar axions produced by Compton and bremsstrahlung processes, in the resonant excitation of  $^{169}\text{Tm}$ , constraining the axion-electron  $\times$  axion-nucleon couplings.
- 4 ANDRIAMONJE 10 search for solar axions produced from  $^7\text{Li}$  (478 keV) and  $\text{D}(p,\gamma)^3\text{He}$  (5.5 MeV) nuclear transitions. They show limits on the axion-photon coupling for two reference values of the axion-nucleon coupling for  $m_A < 100$  eV.
- 5 This is an update of HANNESTAD 08 including 7 years of WMAP data.
- 6 ANDRIAMONJE 09 look for solar axions produced from the thermally excited 14.4 keV level of  $^{57}\text{Fe}$ . They show limits on the axion-nucleon  $\times$  axion-photon coupling assuming  $m_A < 0.03$  eV.
- 7 DERBIN 09A look for Primakoff-produced solar axions in the resonant excitation of  $^{169}\text{Tm}$ , constraining the axion-photon  $\times$  axion-nucleon couplings.
- 8 KEKEZ 09 look at axio-electric effect of solar axions in HPGe detectors. The one-loop axion-electron coupling for hadronic axions is used.
- 9 This is an update of HANNESTAD 07 including 5 years of WMAP data.
- 10 This is an update of HANNESTAD 05A with new cosmological data, notably WMAP (3 years) and baryon acoustic oscillations (BAO). Lyman- $\alpha$  data are left out, in contrast to HANNESTAD 05A and MELCHIORRI 07A, because it is argued that systematic errors are large. It uses Bayesian statistics and marginalizes over a possible neutrino hot dark matter component.
- 11 MELCHIORRI 07A is analogous to HANNESTAD 05A, with updated cosmological data, notably WMAP (3 years). Uses Bayesian statistics and marginalizes over a possible

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

- neutrino hot dark matter component. Leaving out Lyman- $\alpha$  data, a conservative limit is 1.4 eV.
- <sup>12</sup> HANNESTAD 05A puts an upper limit on the mass of hadronic axion because in this mass range it would have been thermalized and contribute to the hot dark matter component of the universe. The limit is based on the CMB anisotropy from WMAP, SDSS large scale structure, Lyman  $\alpha$ , and the prior Hubble parameter from HST Key Project. A  $\chi^2$  statistic is used. Neutrinos are assumed not to contribute to hot dark matter.
- <sup>13</sup> MOROI 98 points out that a KSVZ axion of this mass range (see CHANG 93) can be a viable hot dark matter of Universe, as long as the model-dependent  $g_{A\gamma}$  is accidentally small enough as originally emphasized by KAPLAN 85; see Fig. 1.
- <sup>14</sup> BORISOV 97 bound is on the axion-electron coupling  $g_{ae} < 1 \times 10^{-13}$  from the photo-production of axions off of magnetic fields in the outer layers of neutron stars.
- <sup>15</sup> KACHELRIESS 97 bound is on the axion-electron coupling  $g_{ae} < 1 \times 10^{-10}$  from the production of axions in strongly magnetized neutron stars. The authors also quote a stronger limit,  $g_{ae} < 9 \times 10^{-13}$  which is strongly dependent on the strength of the magnetic field in white dwarfs.
- <sup>16</sup> KEIL 97 uses new measurements of the axial-vector coupling strength of nucleons, as well as a reanalysis of many-body effects and pion-emission processes in the core of the neutron star, to update limits on the invisible-axion mass.
- <sup>17</sup> RAFFELT 95 reexamined the constraints on axion emission from red giants due to the axion-electron coupling. They improve on DEARBORN 86 by taking into proper account degeneracy effects in the bremsstrahlung rate. The limit comes from requiring the red giant core mass at helium ignition not to exceed its standard value by more than 5% (0.025 solar masses).
- <sup>18</sup> ALTHERR 94 bound is on the axion-electron coupling  $g_{ae} < 1.5 \times 10^{-13}$ , from energy loss via axion emission.
- <sup>19</sup> CHANG 93 updates ENGEL 90 bound with the Kaplan-Manohar ambiguity in  $z=m_u/m_d$  (see the Note on the Quark Masses in the Quark Particle Listings). It leaves the window  $f_A=3 \times 10^5 - 3 \times 10^6$  GeV open. The constraint from Big-Bang Nucleosynthesis is satisfied in this window as well.
- <sup>20</sup> BERSHADY 91 searched for a line at wave length from 3100–8300 Å expected from 2 $\gamma$  decays of relic thermal axions in intergalactic light of three rich clusters of galaxies.
- <sup>21</sup> KIM 91C argues that the bound from the mass density of the universe will change drastically for the supersymmetric models due to the entropy production of axion (scalar component in the axionic chiral multiplet) decay. Note that it is an *upperbound* rather than a lowerbound.
- <sup>22</sup> RAFFELT 91B argue that previous SN1987A bounds must be relaxed due to corrections to nucleon bremsstrahlung processes.
- <sup>23</sup> RESSELL 91 uses absence of any intracluster line emission to set limit.
- <sup>24</sup> ENGEL 90 rule out  $10^{-10} \lesssim g_{AN} \lesssim 10^{-3}$ , which for a hadronic axion with EMC motivated axion-nucleon couplings corresponds to  $2.5 \times 10^{-3} \text{ eV} \lesssim m_{A^0} \lesssim 2.5 \times 10^4 \text{ eV}$ . The constraint is loose in the middle of the range, i.e. for  $g_{AN} \sim 10^{-6}$ .
- <sup>25</sup> RAFFELT 90D is a re-analysis of DEARBORN 86.
- <sup>26</sup> The region  $m_{A^0} \gtrsim 2 \text{ eV}$  is also allowed.
- <sup>27</sup> ERICSON 89 considered various nuclear corrections to axion emission in a supernova core, and found a reduction of the previous limit (MAYLE 88) by a large factor.
- <sup>28</sup> MAYLE 89 limit based on naive quark model couplings of axion to nucleons. Limit based on couplings motivated by EMC measurements is 2–4 times weaker. The limit from axion-electron coupling is weak: see HATSUDA 88B.
- <sup>29</sup> RAFFELT 88B derives a limit for the energy generation rate by exotic processes in helium-burning stars  $\epsilon < 100 \text{ erg g}^{-1} \text{ s}^{-1}$ , which gives a firmer basis for the axion limits based on red giant cooling.
- <sup>30</sup> RAFFELT 87 also gives a limit  $g_{A\gamma} < 1 \times 10^{-10} \text{ GeV}^{-1}$ .
- <sup>31</sup> DEARBORN 86 also gives a limit  $g_{A\gamma} < 1.4 \times 10^{-11} \text{ GeV}^{-1}$ .
- <sup>32</sup> RAFFELT 86 gives a limit  $g_{A\gamma} < 1.1 \times 10^{-10} \text{ GeV}^{-1}$  from red giants and  $< 2.4 \times 10^{-9} \text{ GeV}^{-1}$  from the sun.
- <sup>33</sup> KAPLAN 85 says  $m_{A^0} < 23 \text{ eV}$  is allowed for a special choice of model parameters.
- <sup>34</sup> FUKUGITA 82 gives a limit  $g_{A\gamma} < 2.3 \times 10^{-10} \text{ GeV}^{-1}$ .

### Search for Relic Invisible Axions

Limits are for  $[G_{A\gamma\gamma}/m_{A^0}]^2 \rho_A$  where  $G_{A\gamma\gamma}$  denotes the axion two-photon coupling,

$$L_{\text{int}} = -\frac{G_{A\gamma\gamma}}{4} F_{\mu\nu} \tilde{F}^{\mu\nu} = G_{A\gamma\gamma} \mathbf{E} \cdot \mathbf{B}, \text{ and } \rho_A \text{ is the axion energy density near the earth.}$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		1 BECK	13	$m_{A^0} = 0.11 \text{ MeV}$
$< 3.5 \times 10^{-43}$		2 HOSKINS	11	ADMX $m_{A^0} = 3.3\text{--}3.69 \times 10^{-6} \text{ eV}$
$< 2.9 \times 10^{-43}$	90	3 ASZTALOS	10	ADMX $m_{A^0} = 3.34\text{--}3.53 \times 10^{-6} \text{ eV}$
$< 1.9 \times 10^{-43}$	97.7	4 DUFFY	06	ADMX $m_{A^0} = 1.98\text{--}2.17 \times 10^{-6} \text{ eV}$
$< 5.5 \times 10^{-43}$	90	5 ASZTALOS	04	ADMX $m_{A^0} = 1.9\text{--}3.3 \times 10^{-6} \text{ eV}$
		6 KIM	98	THEO
$< 2 \times 10^{-41}$		7 HAGMANN	90	CNTR $m_{A^0} = (5.4\text{--}5.9) \times 10^{-6} \text{ eV}$
$< 1.3 \times 10^{-42}$	95	8 WUENSCH	89	CNTR $m_{A^0} = (4.5\text{--}10.2) \times 10^{-6} \text{ eV}$
$< 2 \times 10^{-41}$	95	8 WUENSCH	89	CNTR $m_{A^0} = (11.3\text{--}16.3) \times 10^{-6} \text{ eV}$

<sup>1</sup> BECK 13 argues that dark-matter axions passing through Earth may generate a small observable signal in resonant S/N/S Josephson junctions. A measurement by HOFFMANN 04 [Physical Review B70 180503 (2004)] is interpreted in terms of subdominant dark matter axions with  $m_{A^0} = 0.11 \text{ MeV}$ .

<sup>2</sup> HOSKINS 11 is analogous to DUFFY 06. See Fig. 4 for the mass-dependent limit in terms of the local density.

<sup>3</sup> ASZTALOS 10 used the upgraded detector of ASZTALOS 04 to search for halo axions. See their Fig. 5 for the  $m_{A^0}$  dependence of the limit.

<sup>4</sup> DUFFY 06 used the upgraded detector of ASZTALOS 04, while assuming a smaller velocity dispersion than the isothermal model as in Eq. (8) of their paper. See Fig. 10 of their paper on the axion mass dependence of the limit.

<sup>5</sup> ASZTALOS 04 looked for a conversion of halo axions to microwave photons in magnetic field. At 90% CL, the KSVZ axion cannot have a local halo density more than

0.45  $\text{GeV}/\text{cm}^3$  in the quoted mass range. See Fig. 7 of their paper on the axion mass dependence of the limit.

<sup>6</sup> KIM 98 calculated the axion-to-photon couplings for various axion models and compared them to the HAGMANN 90 bounds. This analysis demonstrates a strong model dependence of  $G_{A\gamma\gamma}$  and hence the bound from relic axion search.

<sup>7</sup> HAGMANN 90 experiment is based on the proposal of SIKIVIE 83.

<sup>8</sup> WUENSCH 89 looks for condensed axions near the earth that could be converted to photons in the presence of an intense electromagnetic field via the Primakoff effect, following the proposal of SIKIVIE 83. The theoretical prediction with  $[G_{A\gamma\gamma}/m_{A^0}]^2 = 2 \times 10^{-14} \text{ MeV}^{-4}$  (the three generation DFSZ model) and  $\rho_A = 300 \text{ MeV}/\text{cm}^3$  that makes up galactic halos gives  $(G_{A\gamma\gamma}/m_{A^0})^2 \rho_A = 4 \times 10^{-44}$ . Note that our definition of  $G_{A\gamma\gamma}$  is  $(1/4\pi)$  smaller than that of WUENSCH 89.

### Invisible $A^0$ (Axion) Limits from Photon Coupling

Limits are for the modulus of the axion-two-photon coupling  $G_{A\gamma\gamma}$  defined by

$L = -G_{A\gamma\gamma} \phi_A \mathbf{E} \cdot \mathbf{B}$ . For scalars  $S^0$  the limit is on the coupling constant in

$L = G_{S\gamma\gamma} \phi_S (\mathbf{E}^2 - \mathbf{B}^2)$ . The relation between  $G_{A\gamma\gamma}$  and  $m_{A^0}$  is not used unless stated otherwise, i.e., many of these bounds apply to low-mass axion-like particles (ALPs), not to QCD axions.

VALUE ( $\text{GeV}^{-1}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		1 ANASTASSO...	15	CAST Chameleons
$< 1.47 \times 10^{-10}$	95	2 ARIK	15	CAST $m_{A^0} = 0.39\text{--}0.42 \text{ eV}$
$< 3.5 \times 10^{-8}$	95	3 BALLOU	15	LSW $m_{A^0} < 2 \times 10^{-4} \text{ eV}$
		4 BRAX	15	ASTR $m_{S^0} < 4 \times 10^{-12} \text{ eV}$
$< 5.42 \times 10^{-4}$	95	5 HASEBE	15	LASR $m_{A^0} = 0.15 \text{ eV}$
		6 MILLEA	15	COSM Axion-like particles
		7 VANTILBURG	15	Dilaton-like dark matter
$< 4.1 \times 10^{-10}$	99.7	8 VINYOLE	15	ASTR $m_{A^0} = 0.6\text{--}185 \text{ eV}$
$< 3.3 \times 10^{-10}$	95	9 ARIK	14	CAST $m_{A^0} = 0.64\text{--}1.17 \text{ eV}$
$< 6.6 \times 10^{-11}$	95	10 AYALA	14	ASTR Globular clusters
$< 1.4 \times 10^{-7}$	95	11 DELLA-VALLE	14	$m_{A^0} = 1 \text{ MeV}$
		12 EJLLI	14	COSM $m_{A^0} = 2.66\text{--}48.8 \mu\text{eV}$
$< 8 \times 10^{-8}$	95	13 PUGNAT	14	LSW $m_{A^0} < 0.3 \text{ MeV}$
$< 1 \times 10^{-11}$		14 REESMAN	14	ASTR $m_{A^0} < 1 \times 10^{-10} \text{ eV}$
$< 2.1 \times 10^{-11}$	95	15 ABRAMOWSKI13A	IAC	$m_{A^0} = 15\text{--}60 \text{ neV}$
$< 2.15 \times 10^{-9}$	95	16 ARMENGAUD	13	EDEL $m_{A^0} < 200 \text{ eV}$
$< 4.5 \times 10^{-8}$	95	17 BETZ	13	LSW $m_{A^0} = 7.2 \times 10^{-6} \text{ eV}$
$< 8 \times 10^{-11}$		18 FRIEDLAND	13	ASTR Red giants
$> 2 \times 10^{-11}$		19 MEYER	13	ASTR $m_{A^0} < 1 \times 10^{-7} \text{ eV}$
		20 CADAMURO	12	COSM Axion-like particles
$< 2.5 \times 10^{-13}$	95	21 PAYEZ	12	ASTR $m_{A^0} < 4.2 \times 10^{-14} \text{ eV}$
$< 2.3 \times 10^{-10}$	95	22 ARIK	11	CAST $m_{A^0} = 0.39\text{--}0.64 \text{ eV}$
$< 6.5 \times 10^{-8}$	95	23 EHRET	10	ALPS $m_{A^0} < 0.7 \text{ MeV}$
$< 2.4 \times 10^{-9}$	95	24 AHMED	09A	CDMS $m_{A^0} < 100 \text{ eV}$
$< 1.2\text{--}2.8 \times 10^{-10}$	95	25 ARIK	09	CAST $m_{A^0} = 0.02\text{--}0.39 \text{ eV}$
		26 CHOU	09	Chameleons
$< 7 \times 10^{-10}$		27 GONDOLO	09	ASTR $m_{A^0} < \text{few keV}$
$< 1.3 \times 10^{-6}$	95	28 AFANASEV	08	$m_{S^0} < 1 \text{ MeV}$
$< 3.5 \times 10^{-7}$	99.7	29 CHOU	08	$m_{A^0} < 0.5 \text{ MeV}$
$< 1.1 \times 10^{-6}$	99.7	30 FOUCHÉ	08	$m_{A^0} < 1 \text{ MeV}$
$< 5.6\text{--}13.4 \times 10^{-10}$	95	31 INOUE	08	$m_{A^0} = 0.84\text{--}1.00 \text{ eV}$
$< 5 \times 10^{-7}$		32 ZAVATTINI	08	$m_{A^0} < 1 \text{ MeV}$
$< 8.8 \times 10^{-11}$	95	33 ANDRIAMON...	07	CAST $m_{A^0} < 0.02 \text{ eV}$
$< 1.25 \times 10^{-6}$	95	34 ROBILLIARD	07	$m_{A^0} < 1 \text{ MeV}$
$2\text{--}5 \times 10^{-6}$		35 ZAVATTINI	06	$m_{A^0} = 1\text{--}1.5 \text{ MeV}$
$< 1.1 \times 10^{-9}$	95	36 INOUE	02	$m_{A^0} = 0.05\text{--}0.27 \text{ eV}$
$< 2.78 \times 10^{-9}$	95	37 MORALES	02B	$m_{A^0} < 1 \text{ keV}$
$< 1.7 \times 10^{-9}$	90	38 BERNABEI	01B	$m_{A^0} < 100 \text{ eV}$
$< 1.5 \times 10^{-4}$	90	39 ASTIER	00B	NOMD $m_{A^0} < 40 \text{ eV}$
		40 MASSO	00	THEO induced $\gamma$ coupling
$< 2.7 \times 10^{-9}$	95	41 AVIGNONE	98	SLAX $m_{A^0} < 1 \text{ keV}$
$< 6.0 \times 10^{-10}$	95	42 MORIYAMA	98	$m_{A^0} < 0.03 \text{ eV}$
$< 3.6 \times 10^{-7}$	95	43 CAMERON	93	$m_{A^0} < 10^{-3} \text{ eV}$ , optical rotation
$< 6.7 \times 10^{-7}$	95	44 CAMERON	93	$m_{A^0} < 10^{-3} \text{ eV}$ , photon regeneration
$< 3.6 \times 10^{-9}$	99.7	45 LAZARUS	92	$m_{A^0} < 0.03 \text{ eV}$
$< 7.7 \times 10^{-9}$	99.7	45 LAZARUS	92	$m_{A^0} = 0.03\text{--}0.11 \text{ eV}$
$< 7.7 \times 10^{-7}$	99	46 RUOSO	92	$m_{A^0} < 10^{-3} \text{ eV}$
$< 2.5 \times 10^{-6}$		47 SEMERTZIDIS	90	$m_{A^0} < 7 \times 10^{-4} \text{ eV}$

<sup>1</sup> ANASTASSOPOULOS 15 search for solar chameleons with CAST and derived limits on the chameleon coupling to photons and matter. See their Fig. 12 for the exclusion region.

<sup>2</sup> ARIK 15 is analogous to ARIK 09, and search for solar axions for  $m_{A^0}$  around 0.2 and 0.4 eV. See their Figs. 1 and 3 for the mass-dependent limits.



See key on page 601

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

- <sup>3</sup> Based on OSQAR photon regeneration experiment. See their Fig. 6 for mass-dependent limits on scalar and pseudoscalar bosons.
- <sup>4</sup> BRAX 15 derived limits on conformal and disformal couplings of a scalar to photons by searching for a chaotic absorption pattern in the X-ray and UV bands of the Hydra A galaxy cluster and a BL lac object, respectively. See their Fig. 8.
- <sup>5</sup> HASEBE 15 look for an axion via a four-wave mixing process at quasi-parallel colliding laser beams. They also derived limits on a scalar coupling to photons  $G_{S\gamma\gamma} < 2.62 \times 10^{-4} \text{ GeV}^{-1}$  at  $m_{S^0} = 0.15 \text{ eV}$ . See their Figs. 11 and 12 for mass-dependent limits.
- <sup>6</sup> MILLEA 15 is similar to CADAMURO 12, including the Planck data and the latest inferences of primordial deuterium abundance. See their Fig. 3 for mass-dependent limits.
- <sup>7</sup> VANTILBURG 15 look for harmonic variations in the dysprosium transition frequency data, induced by coherent oscillations of the fine-structure constant due to dilaton-like dark matter, and set the limits,  $G_{S\gamma\gamma} < 6 \times 10^{-27} \text{ GeV}^{-1}$  at  $m_{S^0} = 6 \times 10^{-23} \text{ eV}$ . See their Fig. 4 for mass-dependent limits between  $1 \times 10^{-24} < m_{S^0} < 1 \times 10^{-15} \text{ eV}$ .
- <sup>8</sup> VINOLES 15 performed a global fit analysis based on helioseismology and solar neutrino observations. See their Fig. 9.
- <sup>9</sup> ARIK 14 is similar to ARIK 11. See their Fig. 2 for mass-dependent limits.
- <sup>10</sup> AYALA 14 derived the limit from the helium-burning lifetime of horizontal-branch stars based on number counts in globular clusters.
- <sup>11</sup> DELLA-VALLE 14 use the new PVLAS apparatus to set a limit on vacuum magnetic birefringence induced by axion-like particles. See their Fig. 6 for the mass-dependent limits.
- <sup>12</sup> EULLI 14 set limits on a product of primordial magnetic field and the axion mass using CMB distortion induced by resonant axion production from CMB photons. See their Fig. 1 for limits applying specifically to the DFSZ and KSVZ axion models.
- <sup>13</sup> PUGNAT 14 is analogous to EHRET 10. See their Fig. 5 for mass-dependent limits on scalar and pseudoscalar bosons.
- <sup>14</sup> REESMAN 14 derive limits by requiring effects of axion-photon interconversion on gamma-ray spectra from distant blazars to be no larger than errors in the best-fit optical depth based on a certain extragalactic background light model. See their Fig. 5 for mass-dependent limits.
- <sup>15</sup> ABRAMOWSKI 13a look for irregularities in the energy spectrum of the BL Lac object PKS 2155-304 measured by H.E.S.S. The limits depend on assumed magnetic field around the source. See their Fig. 7 for mass-dependent limits.
- <sup>16</sup> ARMENGAUD 13 is analogous to AVIGNONE 98. See Fig. 6 for the limit.
- <sup>17</sup> BETZ 13 performed a microwave-based light shining through the wall experiment. See their Fig. 13 for mass-dependent limits.
- <sup>18</sup> FRIEDLAND 13 derived the limit by considering blue-loop suppression of the evolution of red giants with 7-12 solar masses.
- <sup>19</sup> MEYER 13 attributed to axion-photon oscillations the observed excess of very high-energy  $\gamma$ -rays with respect to predictions based on extragalactic background light models. See their Fig. 4 for mass-dependent lower limits for various magnetic field configurations.
- <sup>20</sup> CADAMURO 12 derived cosmological limits on  $G_{A\gamma\gamma}$  for axion-like particles. See their Fig. 1 for mass-dependent limits.
- <sup>21</sup> PAYEZ 12 derive limits from polarization measurements of quasar light (see their Fig. 3). The limits depend on assumed magnetic field strength in galaxy clusters. The limits depend on assumed magnetic field and electron density in the local galaxy supercluster.
- <sup>22</sup> ARIK 11 search for solar axions using  $^3\text{He}$  buffer gas in CAST, continuing from the  $^4\text{He}$  version of ARIK 09. See Fig. 2 for the exact mass-dependent limits.
- <sup>23</sup> ALPS is a photon regeneration experiment. See their Fig. 4 for mass-dependent limits on scalar and pseudoscalar bosons.
- <sup>24</sup> AHMED 09a is analogous to AVIGNONE 98.
- <sup>25</sup> ARIK 09 is the  $^4\text{He}$  filling version of the CAST axion helioscope in analogy to INOUE 02 and INOUE 08. See their Fig. 7 for mass-dependent limits.
- <sup>26</sup> CHOU 09 use the GammeV apparatus in the afterglow mode to search for chameleons, (pseudo)scalar bosons with a mass depending on the environment. For pseudoscalars they exclude at  $3\sigma$  the range  $2.6 \times 10^{-7} \text{ GeV}^{-1} < G_{A\gamma\gamma} < 4.2 \times 10^{-6} \text{ GeV}^{-1}$  for vacuum  $m_{A^0}$  roughly below  $6 \text{ MeV}$  for density scaling index exceeding 0.8.
- <sup>27</sup> GONDOLO 09 use the all-flavor measured solar neutrino flux to constrain solar interior temperature and thus energy losses.
- <sup>28</sup> LIPSS photon regeneration experiment, assuming scalar particle  $S^0$ . See Fig. 4 for mass-dependent limits.
- <sup>29</sup> CHOU 08 perform a variable-baseline photon regeneration experiment. See their Fig. 3 for mass-dependent limits. Excludes the PVLAS result of ZAVATTINI 06.
- <sup>30</sup> FOUCHÉ 08 is an update of ROBILLIARD 07. See their Fig. 12 for mass-dependent limits.
- <sup>31</sup> INOUE 08 is an extension of INOUE 02 to larger axion masses, using the Tokyo axion helioscope. See their Fig. 4 for mass-dependent limits.
- <sup>32</sup> ZAVATTINI 08 is an upgrade of ZAVATTINI 06, see their Fig. 8 for mass-dependent limits. They now exclude the parameter range where ZAVATTINI 06 had seen a positive signature.
- <sup>33</sup> ANDRIAMONJE 07 looked for Primakoff conversion of solar axions in 9T superconducting magnet into X-rays. Supersedes ZIOUTAS 05.
- <sup>34</sup> ROBILLIARD 07 perform a photon regeneration experiment with a pulsed laser and pulsed magnetic field. See their Fig. 4 for mass-dependent limits. Excludes the PVLAS result of ZAVATTINI 06 with a CL exceeding 99.9%.
- <sup>35</sup> ZAVATTINI 06 propagate a laser beam in a magnetic field and observe dichroism and birefringence effects that could be attributed to an axion-like particle. This result is now excluded by ROBILLIARD 07, ZAVATTINI 08, and CHOU 08.
- <sup>36</sup> INOUE 02 looked for Primakoff conversion of solar axions in 4T superconducting magnet into X-ray.
- <sup>37</sup> MORALES 02b looked for the coherent conversion of solar axions to photons via the Primakoff effect in Germanium detector.
- <sup>38</sup> BERNABE 01b looked for Primakoff coherent conversion of solar axions into photons via Bragg scattering in NaI crystal in DAMA dark matter detector.
- <sup>39</sup> ASTIER 00b looked for production of axions from the interaction of high-energy photons with the horn magnetic field and their subsequent re-conversion to photons via the interaction with the NOMAD dipole magnetic field.
- <sup>40</sup> MASSO 00 studied limits on axion-photon coupling using the induced axion-photon coupling through the proton loop and CAMERON 93 bound on the axion-photon coupling using optical rotation. They obtained the bound  $g_{p\gamma}^2/4\pi < 1.7 \times 10^{-9}$  for the coupling  $g_{p\gamma}^2 P \phi A$ .

- <sup>41</sup> AVIGNONE 98 result is based on the coherent conversion of solar axions to photons via the Primakoff effect in a single crystal germanium detector.
- <sup>42</sup> Based on the conversion of solar axions to X-rays in a strong laboratory magnetic field.
- <sup>43</sup> Experiment based on proposal by MAIANI 86.
- <sup>44</sup> Experiment based on proposal by VANBIBBER 87.
- <sup>45</sup> LAZARUS 92 experiment is based on proposal found in VANBIBBER 89.
- <sup>46</sup> RUOSO 92 experiment is based on the proposal by VANBIBBER 87.
- <sup>47</sup> SEMERTZIDIS 90 experiment is based on the proposal of MAIANI 86. The limit is obtained by taking the noise amplitude as the upper limit. Limits extend to  $m_{A^0} = 4 \times 10^{-3}$  where  $G_{A\gamma\gamma} < 1 \times 10^{-4} \text{ GeV}^{-1}$ .

### Limit on Invisible $A^0$ (Axion) Electron Coupling

The limit is for  $G_{Aee}\partial_\mu\phi A\bar{\psi}\gamma^\mu\psi$  in  $\text{GeV}^{-1}$ , or equivalently, the dipole-dipole potential  $\frac{G_{Aee}^2}{4\pi}((\sigma_1 \cdot \sigma_2) - 3(\sigma_1 \cdot \mathbf{n})(\sigma_2 \cdot \mathbf{n}))/r^3$  where  $\mathbf{n}=\mathbf{r}/r$ .

VALUE ( $\text{GeV}^{-1}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 7.8 \times 10^{-10}$	90	1 ABE	14F	XMAS $m_{A^0} = 60 \text{ keV}$
$< 7.5 \times 10^{-9}$	90	2 APRILE	14B	X100 Solar axions
$< 1 \times 10^{-9}$	90	3 APRILE	14B	X100 $m_{A^0} = 5-7 \text{ keV}$
$< 0.94-8.0 \times 10^{-5}$	90	4 DERBIN	14	CNTR $m_{A^0} = 0.1-1 \text{ MeV}$
$< 3 \times 10^{-10}$	99	5 MILLER-BER...	14	ASTR White dwarf cooling
$< 5.3 \times 10^{-8}$	90	6 ABE	13D	XMAS Solar axions
$< 1.05 \times 10^{-9}$	90	7 ARMENGAUD	13	EDEL $m_{A^0} = 12.5 \text{ keV}$
$< 2.53 \times 10^{-8}$	90	8 ARMENGAUD	13	EDEL Solar axions
$< 1.4-9.5 \times 10^{-4}$	90	9 BARTH	13	CAST Solar axions
$< 2.9 \times 10^{-5}$	68	10 DERBIN	13	CNTR $m_{A^0} = 0.1-1 \text{ MeV}$
$< 4.2 \times 10^{-10}$	95	11 HECKEL	13	$m_{A^0} \leq 0.1 \mu\text{eV}$
$< 7 \times 10^{-10}$	95	12 VIAUX	13A	ASTR Low-mass red giants
$< 2.2 \times 10^{-7}$	90	13 CORSICO	12	ASTR White dwarf cooling
$< 0.02-1 \times 10^{-7}$	90	14 DERBIN	12	CNTR Solar axions
$< 1.4 \times 10^{-9}$	90	15 AALSETH	11	CNTR $m_{A^0} = 0.3-8 \text{ keV}$
$< 3 \times 10^{-6}$	90	16 AHMED	09A	CDMS $m_{A^0} = 2.5 \text{ keV}$
$< 5.3 \times 10^{-5}$	66	17 DAVOUDIASEL	09	ASTR Earth cooling
$< 6.7 \times 10^{-5}$	66	18 NI	94	Induced magnetism
$< 3.6 \times 10^{-4}$	66	19 CHUI	93	Induced magnetism
$< 2.7 \times 10^{-5}$	66	20 PAN	92	Torsion pendulum
$< 1.9 \times 10^{-3}$	66	21 BOBRAKOV	91	Induced magnetism
$< 8.9 \times 10^{-4}$	66	22 WINELAND	91	NMR
$< 6 \times 10^{-5}$	95	23 RITTER	90	Torsion pendulum
	95	24 VOROBYOV	88	Induced magnetism

- <sup>1</sup> ABE 14F set limits on the axioelectric effect in the XMAS detector assuming the pseudoscalar constitutes all the local dark matter. See their Fig. 3 for limits between  $m_{A^0} = 40-120 \text{ keV}$ .
- <sup>2</sup> APRILE 14B look for solar axions using the XENON100 detector.
- <sup>3</sup> APRILE 14B is analogous to AHMED 09A. See their Fig. 7 for limits between  $1 \text{ keV} < m_{A^0} < 35 \text{ keV}$ .
- <sup>4</sup> DERBIN 14 is an update of DERBIN 13 with a BGO scintillating bolometer. See their Fig. 3 for mass-dependent limits.
- <sup>5</sup> MILLER-BERTOLAMI 14 studied the impact of axion emission on white dwarf cooling in a self-consistent way.
- <sup>6</sup> ABE 13D is analogous to DERBIN 12, using the XMAS detector.
- <sup>7</sup> ARMENGAUD 13 is similar to AALSETH 11. See their Fig. 10 for limits between  $3 \text{ keV} < m_{A^0} < 100 \text{ keV}$ .
- <sup>8</sup> ARMENGAUD 13 is similar to DERBIN 12, and take account of axio-recombination and axio-deexcitation effects. See their Fig. 12 for mass-dependent limits.
- <sup>9</sup> BARTH 13 search for solar axions produced by axion-electron coupling, and obtained the limit,  $G_{Aee} \cdot G_{A\gamma\gamma} < 7.9 \times 10^{-20} \text{ GeV}^{-2}$  at 95%CL.
- <sup>10</sup> DERBIN 13 looked for 5.5 MeV solar axions produced in  $p d \rightarrow ^3\text{He} A^0$  in a BGO detector through the axioelectric effect. See their Fig. 4 for mass-dependent limits.
- <sup>11</sup> HECKEL 13 studied the influence of 2 or 4 stationary sources each containing  $6.0 \times 10^{24}$  polarized electrons, on a rotating torsion pendulum containing  $9.8 \times 10^{24}$  polarized electrons. See their Fig. 4 for mass-dependent limits.
- <sup>12</sup> VIAUX 13A constrain axion emission using the observed brightness of the tip of the red-giant branch in the globular cluster M5.
- <sup>13</sup> CORSICO 12 attributed the excessive cooling rate of the pulsating white dwarf R548 to emission of axions with  $G_{Aee} \simeq 5 \times 10^{-10}$ .
- <sup>14</sup> DERBIN 12 look for solar axions with the axio-electric effect in a Si(Li) detector. The solar production is based on Compton and bremsstrahlung processes.
- <sup>15</sup> AALSETH 11 is analogous to AHMED 09A. See their Fig. 4 for mass-dependent limits.
- <sup>16</sup> AHMED 09A assume keV-mass pseudoscalars are the local dark matter and constrain the axio-electric effect in the CDMS detector. See their Fig. 5 for mass-dependent limits.
- <sup>17</sup> DAVOUDIASEL 09 use geophysical constraints on Earth cooling by axion emission.
- <sup>18</sup> These experiments measured induced magnetization of a bulk material by the spin-dependent potential generated from other bulk material with aligned electron spins, where the magnetic field is shielded with superconductor.
- <sup>19</sup> These experiments used a torsion pendulum to measure the potential between two bulk matter objects where the spins are polarized but without a net magnetic field in either of them.
- <sup>20</sup> WINELAND 91 looked for an effect of bulk matter with aligned electron spins on atomic hyperfine splitting using nuclear magnetic resonance.

# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and Other Very Light Bosons

### Invisible $A^0$ (Axion) Limits from Nucleon Coupling

Limits are for the axion mass in eV.

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1 $\times 10^2$	95	1 GAVRILYUK 15	CNTR	Solar axion
		2 KLIMCHITSK...15		Casimir-less
		3 BEZERRA 14		Casimir effect
		4 BEZERRA 14A		Casimir effect
		5 BEZERRA 14B		Casimir effect
		6 BEZERRA 14C		Casimir effect
		7 BLUM 14	COSM	$^4\text{He}$ abundance
		8 LEINSON 14	ASTR	Neutron star cooling
<2.50 $\times 10^2$	95	9 ALESSANDRIA 13	CNTR	Solar axion
		10 ARMENGAUD 13	EDEL	Solar axion
<1.55 $\times 10^2$	90	11 BELLI 12	CNTR	Solar axion
<8.6 $\times 10^3$	90	12 BELLINI 12B	BORX	Solar axion
<1.41 $\times 10^2$	90	13 DERBIN 11	CNTR	Solar axion
<1.45 $\times 10^2$	95	14 BELLINI 08	CNTR	Solar axion
		15 ADELBERGER 07		Test of Newton's law

- <sup>1</sup> GAVRILYUK 15 look for solar axions emitted by the M1 transition of  $^{83}\text{Kr}$  (9.4 keV). The mass bound assumes  $m_U/m_d = 0.56$  and  $S = 0.5$ .
- <sup>2</sup> KLIMCHITSKAYA 15 use the measurement of differential forces between a test mass and rotating source masses of Au and Si to constrain the force due to two-axion exchange for  $1.7 \times 10^{-3} < m_{A^0} < 0.9$  eV. See their Figs. 1 and 2 for mass dependent limits.
- <sup>3</sup> BEZERRA 14 use the measurement of the thermal Casimir-Polder force between a Bose-Einstein condensate of  $^{87}\text{Rb}$  atoms and a  $\text{SiO}_2$  plate to constrain the force mediated by exchange of two pseudoscalars for  $0.1 \text{ meV} < m_{A^0} < 0.3$  eV. See their Fig. 2 for the mass-dependent limit on pseudoscalar coupling to nucleons.
- <sup>4</sup> BEZERRA 14A is analogous to BEZERRA 14. They use the measurement of the Casimir pressure between two Au-coated plates to constrain pseudoscalar coupling to nucleons for  $1 \times 10^{-3} \text{ eV} < m_{A^0} < 15$  eV. See their Figs. 1 and 2 for the mass-dependent limit.
- <sup>5</sup> BEZERRA 14B is analogous to BEZERRA 14. BEZERRA 14B use the measurement of the normal and lateral Casimir forces between sinusoidally corrugated surfaces of a sphere and a plate to constrain pseudoscalar coupling to nucleons for  $1 \text{ eV} < m_{A^0} < 20$  eV. See their Figs. 1–3 for mass-dependent limits.
- <sup>6</sup> BEZERRA 14C is analogous to BEZERRA 14. They use the measurement of the gradient of the Casimir force between Au- and Ni-coated surfaces of a sphere and a plate to constrain pseudoscalar coupling to nucleons for  $3 \times 10^{-5} \text{ eV} < m_{A^0} < 1$  eV. See their Figs. 1, 3, and 4 for the mass-dependent limits.
- <sup>7</sup> BLUM 14 studied effects of an oscillating strong  $CP$  phase induced by axion dark matter on the primordial  $^4\text{He}$  abundance. See their Fig. 1 for mass-dependent limits.
- <sup>8</sup> LEINSON 14 attributes the excessive cooling rate of the neutron star in Cassiopeia A to axion emission from the superfluid core, and found  $C_n^2 m_n^2 A^0 \simeq 5.7 \times 10^{-6} \text{ eV}^2$ , where  $C_n$  is the effective Peccei-Quinn charge of the neutron.
- <sup>9</sup> ALESSANDRIA 13 used the CUORE experiment to look for 14.4 keV solar axions produced from the M1 transition of thermally excited  $^{57}\text{Fe}$  nuclei in the solar core, using the axio-electric effect. The limit assumes the hadronic axion model. See their Fig. 4 for the limit on product of axion couplings to electrons and nucleons.
- <sup>10</sup> ARMENGAUD 13 is analogous to ALESSANDRIA 13. The limit assumes the hadronic axion model. See their Fig. 8 for the limit on product of axion couplings to electrons and nucleons.
- <sup>11</sup> BELLI 12 looked for solar axions emitted by the M1 transition of  $^7\text{Li}^*$  (478 keV) after the electron capture of  $^7\text{Be}$ , using the resonant excitation  $^7\text{Li}$  in the LiF crystal. The mass bound assumes  $m_U/m_d = 0.55$ ,  $m_U/m_S = 0.029$ , and the flavor-singlet axial vector matrix element  $S = 0.4$ .
- <sup>12</sup> BELLINI 12B looked for 5.5 MeV solar axions produced in the  $pd \rightarrow ^3\text{He} A^0$ . The limit assumes the hadronic axion model. See their Figs. 4 and 5 for mass-dependent limits on products of axion couplings to photons, electrons, and nucleons.
- <sup>13</sup> DERBIN 11 looked for solar axions emitted by the M1 transition of thermally excited  $^{57}\text{Fe}$  nuclei in the Sun, using their possible resonant capture on  $^{57}\text{Fe}$  in the laboratory. The mass bound assumes  $m_U/m_d = 0.56$  and the flavor-singlet axial vector matrix element  $S = 3F - D \simeq 0.5$ .
- <sup>14</sup> BELLINI 08 consider solar axions emitted in the M1 transition of  $^7\text{Li}^*$  (478 keV) and look for a peak at 478 keV in the energy spectra of the Counting Test Facility (CTF), a Borexino prototype. For  $m_{A^0} < 450$  keV they find mass-dependent limits on products of axion couplings to photons, electrons, and nucleons.
- <sup>15</sup> ADELBERGER 07 use precision tests of Newton's law to constrain a force contribution from the exchange of two pseudoscalars. See their Fig. 5 for limits on the pseudoscalar coupling to nucleons, relevant for  $m_{A^0}$  below about 1 MeV.

### Axion Limits from $T$ -violating Medium-Range Forces

The limit is for the coupling  $g = g_p g_s$  in a  $T$ -violating potential between nucleons or nucleon and electron of the form  $V = \frac{g\hbar^2}{8\pi m_p} (\boldsymbol{\sigma} \cdot \hat{\mathbf{r}}) \left( \frac{1}{r^2} + \frac{1}{\lambda r} \right) e^{-r/\lambda}$ , where  $g_p$  and  $g_s$  are dimensionless scalar and pseudoscalar coupling constants and  $\lambda = \hbar/(m_A c)$  is the range of the force.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<1.7 $\times 10^{-6}$	1 AFACH 15		ultracold neutrons
	2 STADNIK 15	THEO	nucleon spin contributions for nuclei
	3 BULATOWICZ 13	NMR	polarized $^{129}\text{Xe}$ and $^{131}\text{Xe}$
	4 CHU 13		polarized $^3\text{He}$

<sup>5</sup> TULLNEY 13	SQID	polarized $^3\text{He}$ and $^{129}\text{Xe}$
<sup>6</sup> RAFFELT 12		stellar energy loss
<sup>7</sup> HOEDL 11		torsion pendulum
<sup>8</sup> PETUKHOV 10		polarized $^3\text{He}$
<sup>9</sup> SEREBROV 10		ultracold neutrons
<sup>10</sup> IGNATOVICH 09	RVUE	ultracold neutrons
<sup>11</sup> SEREBROV 09	RVUE	ultracold neutrons
<sup>12</sup> BAESSLER 07		ultracold neutrons
<sup>13</sup> HECKEL 06		torsion pendulum
<sup>14</sup> NI 99		paramagnetic Tb F <sub>3</sub>
<sup>15</sup> POSPELOV 98	THEO	neutron EDM
<sup>16</sup> YODIN 96		
<sup>17</sup> RITTER 93		torsion pendulum
<sup>18</sup> VENEMA 92		nuclear spin-precession frequencies
<sup>19</sup> WINELAND 91	NMR	

- <sup>1</sup> AFACH 15 look for a change of spin precession frequency of ultracold neutrons when a magnetic field with opposite directions is applied, and find  $g < 2.2 \times 10^{-27} (\text{m}/\lambda)^2$  at 95% CL for  $1 \mu\text{m} < \lambda < 5$  mm. See their Fig. 3 for their limits.
- <sup>2</sup> STADNIK 15 studied proton and neutron spin contributions for nuclei and derive the limits  $g < 10^{-28} - 10^{-23}$  for  $\lambda > 3 \times 10^{-4}$  m using the data of TULLNEY 13. See their Figs. 1 and 2 for  $\lambda$ -dependent limits.
- <sup>3</sup> BULATOWICZ 13 looked for NMR frequency shifts in polarized  $^{129}\text{Xe}$  and  $^{131}\text{Xe}$  when a zirconia rod is positioned near the NMR cell, and find  $g < 1 \times 10^{-19} - 1 \times 10^{-24}$  for  $\lambda = 0.01$ –1 cm. See their Fig. 4 for their limits.
- <sup>4</sup> CHU 13 look for a shift of the spin precession frequency of polarized  $^3\text{He}$  in the presence of an unpolarized mass, in analogy to YODIN 96. See Fig. 3 for limits on  $g$  in the approximate  $m_{A^0}$  range 0.02–2 meV.
- <sup>5</sup> TULLNEY 13 look for a shift of the precession frequency difference between the colocated  $^3\text{He}$  and  $^{129}\text{Xe}$  in the presence an unpolarized mass, and derive limits  $g < 3 \times 10^{-29} - 2 \times 10^{-22}$  for  $\lambda > 3 \times 10^{-4}$  m. See their Fig. 3 for  $\lambda$ -dependent limits.
- <sup>6</sup> RAFFELT 12 show that the pseudoscalar couplings to electron and nucleon and the scalar coupling to nucleon are individually constrained by stellar energy-loss arguments and searches for anomalous monopole-monopole forces, together providing restrictive constraints on  $g$ . See their Figs. 2 and 3 for results.
- <sup>7</sup> HOEDL 11 use a novel torsion pendulum to study the force by the polarized electrons of an external magnet. In their Fig. 3 they show restrictive limits on  $g$  in the approximate  $m_{A^0}$  range 0.03–10 meV.
- <sup>8</sup> PETUKHOV 10 use spin relaxation of polarized  $^3\text{He}$  and find  $g < 3 \times 10^{-23} (\text{cm}/\lambda)^2$  at 95% CL for the force range  $\lambda = 10^{-4}$ –1 cm.
- <sup>9</sup> SEREBROV 10 use spin precession of ultracold neutrons close to bulk matter and find  $g < 2 \times 10^{-21} (\text{cm}/\lambda)^2$  at 95% CL for the force range  $\lambda = 10^{-4}$ –1 cm.
- <sup>10</sup> IGNATOVICH 09 use data on depolarization of ultracold neutrons in material traps. They show  $\lambda$ -dependent limits in their Fig. 1.
- <sup>11</sup> SEREBROV 09 uses data on depolarization of ultracold neutrons stored in material traps and finds  $g < 2.96 \times 10^{-21} (\text{cm}/\lambda)^2$  for the force range  $\lambda = 10^{-3}$ –1 cm and  $g < 3.9 \times 10^{-22} (\text{cm}/\lambda)^2$  for  $\lambda = 10^{-4}$ – $10^{-3}$  cm, each time at 95% CL, significantly improving on BAESSLER 07.
- <sup>12</sup> BAESSLER 07 use the observation of quantum states of ultracold neutrons in the Earth's gravitational field to constrain  $g$  for an interaction range  $1 \mu\text{m}$ –a few mm. See their Fig. 3 for results.
- <sup>13</sup> HECKEL 06 studied the influence of unpolarized bulk matter, including the laboratory's surroundings or the Sun, on a torsion pendulum containing about  $9 \times 10^{22}$  polarized electrons. See their Fig. 4 for limits on  $g$  as a function of interaction range.
- <sup>14</sup> NI 99 searched for a  $T$ -violating medium-range force acting on paramagnetic Tb F<sub>3</sub> salt. See their Fig. 1 for the result.
- <sup>15</sup> POSPELOV 98 studied the possible contribution of  $T$ -violating Medium-Range Force to the neutron electric dipole moment, which is possible when axion interactions violate  $CP$ . The size of the force among nucleons must be smaller than gravity by a factor of  $2 \times 10^{-10} (1 \text{ cm}/\lambda_A)$ , where  $\lambda_A = \hbar/m_A c$ .
- <sup>16</sup> YODIN 96 compared the precession frequencies of atomic  $^{199}\text{Hg}$  and Cs when a large mass is positioned near the cells, relative to an applied magnetic field. See Fig. 3 for their limits.
- <sup>17</sup> RITTER 93 studied the influence of bulk mass with polarized electrons on an unpolarized torsion pendulum, providing limits in the interaction range from 1 to 100 cm.
- <sup>18</sup> VENEMA 92 looked for an effect of Earth's gravity on nuclear spin-precession frequencies of  $^{199}\text{Hg}$  and  $^{201}\text{Hg}$  atoms.
- <sup>19</sup> WINELAND 91 looked for an effect of bulk matter with aligned electron spins on atomic hyperfine resonances in stored  $^9\text{Be}^+$  ions using nuclear magnetic resonance.

### Hidden Photons: Kinetic Mixing Parameter Limits

Hidden photons limits are listed for the first time, including only the most recent papers. Suggestions for previous important results are welcome. Limits are on the kinetic mixing parameter  $\chi$  which is defined by the Lagrangian

$$L = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} - \frac{\chi}{2} F_{\mu\nu} F'^{\mu\nu} + \frac{m^2}{2} A'_\mu A'^\mu,$$

where  $A_\mu$  and  $A'_\mu$  are the photon and hidden-photon fields with field strengths  $F_{\mu\nu}$  and  $F'_{\mu\nu}$ , respectively, and  $m_{\gamma'}$  is the hidden-photon mass.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.7 $\times 10^{-6}$	95	1 KHACHATRY...16	CMS	$m_{\gamma'} = 2$ GeV
<4 $\times 10^{-2}$	95	2 AAD 15CD	ATLS	$m_{\gamma'} = 15$ –55 GeV
		3 ADARE 15		$m_{\gamma'} = 30$ –90 MeV
<1.4 $\times 10^{-3}$	90	4 AN 15A		$m_{\gamma'} = 12$ eV - 40 keV
		5 ANASTASI 15	KLOE	$m_{\gamma'} = 2m_\mu - 1$ GeV
		6 ANASTASI 15A	KLOE	$m_{\gamma'} = 5$ –320 MeV

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# Gauge & Higgs Boson Particle Listings

## Axions ( $A^0$ ) and other Very Light Bosons

$<4.2 \times 10^{-4}$	90	7	BATLEY	15A	NA48	$m_{\gamma'} = 36$ MeV
		8	JAEGLER	15	BELL	$m_{\gamma'} = 0.1\text{--}3.5$ GeV
$<3 \times 10^{-13}$		9	KAZANAS	15	ASTR	$m_{\gamma'} = 2m_e - 100$ MeV
		10	SUZUKI	15		$m_{\gamma'} = 1.9\text{--}4.3$ eV
$<2.3 \times 10^{-13}$	99.7	11	VINYOLES	15	ASTR	$m_{\gamma'} = 8$ eV
		12	ABE	14F	XMAS	$m_{\gamma'} = 40\text{--}120$ keV
$<1.8 \times 10^{-3}$	90	13	AGAKISHIEV	14	HDES	$m_{\gamma'} = 63$ MeV
$<9.0 \times 10^{-4}$	90	14	BABUSCI	14	KLOE	$m_{\gamma'} = 969$ MeV
		15	BATELL	14	BDMP	$m_{\gamma'} = 10^{-3}\text{--}1$ GeV
$<1.3 \times 10^{-7}$	95	16	BLUEMLEIN	14	BDMP	$m_{\gamma'} = 0.6$ GeV
$<3 \times 10^{-18}$		17	FRADETTE	14	COSM	$m_{\gamma'} = 50\text{--}300$ MeV
$<3.5 \times 10^{-4}$	90	18	LEES	14J	BABR	$m_{\gamma'} = 0.2$ GeV
$<9 \times 10^{-4}$	95	19	MERKEL	14	A1	$m_{\gamma'} = 40\text{--}300$ MeV
$<3 \times 10^{-15}$		20	AN	13B	ASTR	$m_{\gamma'} = 2$ keV
$<7 \times 10^{-14}$		21	AN	13C	XE10	$m_{\gamma'} = 100$ eV
$<2.2 \times 10^{-13}$		22	HORVAT	13	HPGE	$m_{\gamma'} = 230$ eV
$<8.06 \times 10^{-5}$	95	23	INADA	13	LSW	$m_{\gamma'} = 0.04$ eV–26 keV
$<2 \times 10^{-10}$	95	24	MIZUMOTO	13		$m_{\gamma'} = 1$ eV
$<1.7 \times 10^{-7}$		25	PARKER	13	LSW	$m_{\gamma'} = 53$ $\mu$ eV
$<5.32 \times 10^{-15}$		26	PARKER	13		$m_{\gamma'} = 53$ $\mu$ eV
$<1 \times 10^{-15}$		27	REDONDO	13	ASTR	$m_{\gamma'} = 2$ keV
$<9 \times 10^{-8}$	95	28	BLUEMLEIN	11	BDMP	$m_{\gamma'} = 70$ MeV

- 1 KHACHATRYAN 16 look for  $\gamma' \rightarrow \mu^+ \mu^-$  in a dark SUSY scenario where the SM-like Higgs boson decays into a pair of the visible lightest neutralinos with mass 10 GeV, both of which decay into  $\gamma'$  and a hidden neutralino with mass 1 GeV. See the right panel in their Fig. 2.
- 2 AAD 15CD look for  $H \rightarrow Z\gamma' \rightarrow 4\ell$  with the ATLAS detector at LHC and find  $\chi < 4\text{--}17 \times 10^{-2}$  for  $m_{\gamma'} = 15\text{--}55$  GeV. See their Fig. 6.
- 3 ADARE 15 look for a hidden photon in  $\pi^0, \eta^0 \rightarrow \gamma e^+ e^-$  at the PHENIX experiment. See their Fig. 4 for mass-dependent limits.
- 4 AN 15A derived limits from the absence of ionization signals in the XENON10 and XENON100 experiments, assuming hidden photons constitute all the local dark matter. Their best limit is  $\chi < 1.3 \times 10^{-15}$  at  $m_{\gamma'} = 18$  eV. See their Fig. 1 for mass-dependent limits.
- 5 ANASTASI 15 look for a production of a hidden photon and a hidden Higgs boson with the KLOE detector at DAΦNE, where the hidden photon decays into a pair of muons and the hidden Higgs boson lighter than  $m_{\gamma'}$  escape detection. See their Figs. 6 and 7 for mass-dependent limits on a product of the hidden fine structure constant and the kinetic mixing.
- 6 ANASTASI 15A look for the decay  $\gamma' \rightarrow e^+ e^-$  in the reaction  $e^+ e^- \rightarrow e^+ e^- \gamma'$ . Limits between  $1.7 \times 10^{-3}$  and  $1 \times 10^{-2}$  are obtained for  $m_{\gamma'} = 5\text{--}320$  MeV (see their Fig. 7).
- 7 BATLEY 15A look for  $\pi^0 \rightarrow \gamma\gamma' (\gamma' \rightarrow e^+ e^-)$  at the NA48/2 experiment. Limits between  $4.2 \times 10^{-4}$  and  $8.8 \times 10^{-3}$  are obtained for  $m_{\gamma'} = 9\text{--}120$  MeV (see their Fig. 4).
- 8 JAEGLER 15 look for the decay  $\gamma' \rightarrow e^+ e^-, \mu^+ \mu^-,$  or  $\pi^+ \pi^-$  in the dark Higgs-gating channel,  $e^+ e^- \rightarrow \gamma' H' (H' \rightarrow \gamma' \gamma')$  at the BELLE experiment. They set limits on a product of the branching fraction and the Born cross section as well as a product of the hidden fine structure constant and the kinetic mixing. See their Figs. 3 and 4.
- 9 KAZANAS 15 set limits by studying the decay of hidden photons  $\gamma' \rightarrow e^+ e^-$  inside and near the progenitor star of SN1987A. See their Fig. 6 for mass-dependent limits.
- 10 SUZUKI 15 looked for hidden-photon dark matter with a dish antenna and derived limits assuming they constitute all the local dark matter. Their limits are  $\chi < 6 \times 10^{-12}$  for  $m_{\gamma'} = 1.9\text{--}4.3$  eV. See their Fig. 7 for mass-dependent limits.
- 11 VINYOLES 15 performed a global fit analysis based on helioseismology and solar neutrino observations, and set the limits  $\chi m_{\gamma'} < 1.8 \times 10^{-12}$  eV for  $m_{\gamma'} = 3 \times 10^{-5}\text{--}8$  eV. See their Fig. 11.
- 12 ABE 14F look for the photoelectric-like interaction in the XMASS detector assuming the hidden photon constitutes all the local dark matter. Limits between  $2 \times 10^{-13}$  and  $1 \times 10^{-12}$  are obtained. See their Fig. 3 for mass-dependent limits.
- 13 AGAKISHIEV 14 look for hidden photons  $\gamma' \rightarrow e^+ e^-$  at the HADES experiment, and set limits on  $\chi$  for  $m_{\gamma'} = 0.02\text{--}0.6$  GeV. See their Fig. 5 for mass-dependent limits.
- 14 BABUSCI 14 look for the decay  $\gamma' \rightarrow \mu^+ \mu^-$  in the reaction  $e^+ e^- \rightarrow \mu^+ \mu^- \gamma'$ . Limits between  $4 \times 10^{-3}$  and  $9.0 \times 10^{-4}$  are obtained for  $520 \text{ MeV} < m_{\gamma'} < 980 \text{ MeV}$  (see their Fig. 7).
- 15 BATELL 14 derived limits from the electron beam dump experiment at SLAC (E-137) by searching for events with recoil electrons by sub-GeV dark matter produced from the decay of the hidden photon. Limits at the level of  $10^{-4}\text{--}10^{-1}$  are obtained for  $m_{\gamma'} = 10^{-3}\text{--}1$  GeV, depending on the dark matter mass and the hidden gauge coupling (see their Fig. 2).
- 16 BLUEMLEIN 14 analyzed the beam dump data taken at the U-70 accelerator to look for  $\gamma'$ -bremsstrahlung and the subsequent decay into muon pairs and hadrons. See their Fig. 4 for mass-dependent excluded regions.
- 17 FRADETTE 14 studied effects of decay of relic hidden photons on BBN and CMB to set constraints on very small values of the kinetic mixing. See their Figs. 4 and 7 for mass-dependent excluded regions.

- 18 LEES 14J look for hidden photons in the reaction  $e^+ e^- \rightarrow \gamma\gamma' (\gamma' \rightarrow e^+ e^-, \mu^+ \mu^-)$ . Limits at the level of  $10^{-4}\text{--}10^{-3}$  are obtained for  $0.02 \text{ GeV} < m_{\gamma'} < 10.2 \text{ GeV}$ . See their Fig. 4 for mass-dependent limits.
- 19 MERKEL 14 look for  $\gamma' \rightarrow e^+ e^-$  at the A1 experiment at the Mainz Microtron (MAMI). See their Fig. 3 for mass-dependent limits.
- 20 AN 13B examined the stellar production of hidden photons, correcting an important error of the production rate of the longitudinal mode which now dominates. See their Fig. 2 for mass-dependent limits based on solar energy loss.
- 21 AN 13C use the solar flux of hidden photons to set a limit on the atomic ionization rate in the XENON10 experiment. They find  $\chi m_{\gamma'} < 3 \times 10^{-12}$  eV for  $m_{\gamma'} < 1$  eV. See their Fig. 2 for mass-dependent limits.
- 22 HORVAT 13 look for hidden-photo-electric effect in HPGe detectors induced by solar hidden photons. See their Fig. 3 for mass-dependent limits.
- 23 INADA 13 search for hidden photons using an intense X-ray beamline at SPring-8. See their Fig. 4 for mass-dependent limits.
- 24 MIZUMOTO 13 look for solar hidden photons. See their Fig. 5 for mass-dependent limits.
- 25 PARKER 13 look for hidden photons using a cryogenic resonant microwave cavity. See their Fig. 5 for mass-dependent limits.
- 26 PARKER 13 derived a limit for the hidden photon CDM with a randomly oriented hidden photon field.
- 27 REDONDO 13 examined the solar emission of hidden photons including the enhancement factor for the longitudinal mode pointed out by AN 13B, and also updated stellar-energy loss arguments. See their Fig. 3 for mass-dependent limits, including a review of the currently best limits from other arguments.
- 28 BLUEMLEIN 11 analyzed the beam dump data taken at the U-70 accelerator to look for  $\pi^0 \rightarrow \gamma\gamma' (\gamma' \rightarrow e^+ e^-)$ . See their Fig. 5 for mass-dependent limits.

## REFERENCES For Searches for Axions ( $A^0$ ) and Other Very Light Bosons

KHACHATRYAN...	16	PL B752 146	V. Khachatryan et al.	(CMS Collab.)
AAD	15CD	PR D92 092001	G. Aad et al.	(ATLAS Collab.)
AAJ	15AZ	PRL 115 161802	R. Aaij et al.	(LHCb Collab.)
ADARE	15	PR C91 031901	A. Adare et al.	(PHENIX Collab.)
AFACH	15	PL B745 58	S. Afach et al.	(ETH, PSI, CAEN, +)
AGOSTINI	15A	EPJ C75 416	M. Agostini et al.	(GERDA Collab.)
AN	15A	PL B747 331	H. An et al.	(CIT, VICT, VIEN)
ANASTASI	15	PL B747 365	A. Anastasi et al.	(KLOE-2 Collab.)
ANASTASI	15A	PL B750 633	A. Anastasi et al.	(KLOE-2 Collab.)
ANASTASSO...	15	PL B749 172	V. Anastassopoulos et al.	(CAST Collab.)
ARIK	15	PR D92 021101	M. Arık et al.	(CAST Collab.)
ARNOLD	15	PR D92 072011	R. Arnold et al.	(NEMO-3 Collab.)
BALLOU	15	PR D92 092002	R. Ballou et al.	(OSQAR Collab.)
BATLEY	15A	PL B746 178	J.R. Batley et al.	(NA48/2 Collab.)
BAYES	15	PR D91 052020	R. Bayes et al.	(TWIST Collab.)
BRAX	15	PR D92 083501	P. Brax, P. Brun, D. Wouters	(SACL, SACL5)
GAVRILYUK	15	JETPL 101 664	Yu.M. Gavrilyuk et al.	
HASEBE	15	PTEP 2015 073C01	T. Hasebe et al.	
JAEGLER	15	PRL 114 211801	I. Jaegle et al.	(BELLE Collab.)
KAZANAS	15	NP B890 17	D. Kazanas et al.	
KLIMCHITSK...	15	EPJ C75 164	G.L. Klimchitskaya, V.M. Mostepanenko	
MILLEA	15	PR D92 023010	M. Millea, L. Knox, B. Fields	(UCD, ILL)
STADNIK	15	EPJ C75 110	Y.V. Stadnik, V.V. Flambaum	(SYDN)
SUZUKI	15	JCAP 1509 042	K. Suzuki et al.	
VANTILBURG	15	PRL 115 011802	K. Van Tilburg et al.	
VINYOLES	15	JCAP 1510 015	N. Vinyoles et al.	
ABE	14F	PRL 113 121301	K. Abe et al.	(XMASS Collab.)
AGAKISHIEV	14	PL B731 265	G. Agakishiev et al.	(HADES Collab.)
ALBERT	14B	PR D90 092004	J.B. Albert et al.	(EXO-200 Collab.)
APRILE	14A	PR D90 062009	E. Aprile et al.	(XENON100 Collab.)
ARIK	14	PRL 112 091302	M. Arık et al.	(CAST Collab.)
AYALA	14	PRL 113 191302	A. Ayala et al.	
BABUSCI	14	PL B736 459	D. Babusci et al.	(KLOE-2 Collab.)
BATELL	14	PRL 113 171802	B. Batell, R. Essig, Z. Surujon	(EFI, STON)
BEZERRA	14	PR D89 035010	V.B. Bezerra et al.	
BEZERRA	14A	EPJ C74 2859	V.B. Bezerra et al.	
BEZERRA	14B	PR D90 055013	V.B. Bezerra et al.	
BEZERRA	14C	PR D89 075002	V.B. Bezerra et al.	
BLUEMLEIN	14	PL B731 320	J. Bluemlein, J. Brunner	(CPPM, DESY)
BLUM	14	PL B737 30	K. Blum et al.	(IAS, PRIN)
DELLA-VALLE	14	PR D90 092003	F. Della Valle et al.	(PVLAS Collab.)
DERBIN	14	EPJ C74 3035	A.V. Derbin et al.	
EJLLI	14	PR D90 123527	D. Ejlli	
FRADETTE	14	PR D90 035022	A. Fradette et al.	
LEES	14J	PRL 113 201801	J.P. Lees et al.	(BABAR Collab.)
LEINSON	14	JCAP 1408 031	L. Leinson	
MERKEL	14	PRL 112 221802	H. Merkel et al.	(A1 at MAMI)
MILLER-BER...	14	JCAP 1410 069	M.M. Miller-Bertolami et al.	
PUGNAT	14	EPJ C74 3027	P. Pugnât et al.	(OSQAR Collab.)
REESMAN	14	JCAP 1408 021	R. Reesman et al.	(OSU)
ABE	13D	PL B724 46	K. Abe et al.	(XMASS Collab.)
ABRAMOWSKI	13A	PR D88 102003	A. Abramowski et al.	(H.E.S.S. Collab.)
ADLARSON	13	PL B726 187	P. Adlarson et al.	(WASA-at-COSY Collab.)
ALESSANDRIA	13	JCAP 1305 007	F. Alessandria et al.	(CUORE Collab.)
AN	13C	PL B725 190	H. An, M. Pospelov, J. Pradler	
AN	13B	PRL 111 041302	H. An, M. Pospelov, J. Pradler	
ARCHIDIACO...	13A	JCAP 1310 020	A. Archidiaco et al.	
ARMENGAUD	13	JCAP 1311 067	E. Armengaud et al.	(EDELWEISS-II Collab.)
BABUSCI	13B	PL B720 111	D. Babusci et al.	(KLOE-2 Collab.)
BARTH	13	JCAP 1305 010	K. Barth et al.	(CAST Collab.)
BECK	13	PRL 111 231801	C. Beck	
BETZ	13	PR D88 075014	M. Betz et al.	(CROWS Collab.)
BULATOWICZ	13	PRL 111 102001	M. Bulatowicz et al.	
CHU	13	PR D87 011105	P.-H. Chu et al.	(DUKE, IND, SJTU)
DERBIN	13	EPJ C73 2490	A. V. Derbin et al.	
FRIEDLAND	13	PRL 110 061101	A. Friedland, M. Giannotti, M. Wise	
HECKEL	13	PRL 111 151802	B. R. Heckel et al.	
HORVAT	13	PL B721 220	R. Horvat et al.	
INADA	13	PL B722 301	T. Inada et al.	
LATTANZI	13	PR D88 063528	M. Lattanzi et al.	
MEYER	13	PR D87 035027	M. Meyer, D. Horns, M. Raue	
MIZUMOTO	13	JCAP 1307 013	T. Mizumoto et al.	
PARKER	13	PR D88 112004	S. Parker et al.	
REDONDO	13	JCAP 1308 034	J. Redondo, G. Raffelt	
TULLNEY	13	PRL 111 100801	K. Tullney et al.	
VIAUX	13A	PRL 111 231301	N. Viaux et al.	
ARCHILLI	12	PL B706 251	F. Archilli et al.	(KLOE-2 Collab.)
BELLI	12	PL B711 41	P. Belli et al.	(DAMA-KIEV)
BELLINI	12B	PR D85 092003	G. Bellini et al.	(Borexino Collab.)
CADAMURO	12	JCAP 1202 032	D. Cadamuro et al.	(MIPM)
CORSICO	12	JCAP 1212 010	A.H. Corsico et al.	(LAPL, RGSUL, WASH+)



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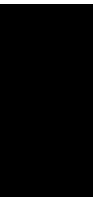
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KORENCHEN...	87	SJNP 46 192	S.M. Korenchenko <i>et al.</i>	(JINR)
MAIER	87	ZPHY A326 527	K. Maier <i>et al.</i>	(STUT, GSI)
MILLS	87	PR D36 707	A.P. Mills, J. Levy	(BELL)
RAFFELT	87	PR D36 2211	G.G. Raffelt, D.S.P. Dearborn	(LLL, UCB)
RIORDAN	87	PRL 59 755	E.M. Riordan <i>et al.</i>	(ROCH, CIT+)
TURNER	87	PRL 59 2489	M.S. Turner	(FNAL, EFI)
VANBIBBER	87	PRL 59 759	K. van Bibber <i>et al.</i>	(LLL, CIT, MIT+)
VONWIMMER.	87	PRL 59 266	U. von Wimmersperg <i>et al.</i>	(WITW)
BADIER	86	ZPHY C31 21	J. Badier <i>et al.</i>	(NA3 Collab.)
BROWN	86	PRL 57 2101	C.N. Brown <i>et al.</i>	(FNAL, WASH, KYOT+)
BRYMAN	86B	PRL 57 2787	D.A. Bryman, E.T.H. Clifford	(TRIUMF)
DAVIER	86	PL B180 295	M. Davier, J. Jeanjean, H. Nguyen Ngoc	(LALO)
DEARBORN	86	PRL 56 26	D.S.P. Dearborn, D.N. Schramm, G. Steigman	(LLL+)
EICHLER	86	PL B175 101	R.A. Eichler <i>et al.</i>	(SINDRUM Collab.)
HALLIN	86	PRL 57 2105	A.L. Hallin <i>et al.</i>	(PRIN)
JODIDIO	86	PR D34 1967	A. Jodidio <i>et al.</i>	(LBL, NWES, TRIU)
Also		PR D37 237 (erratum)	A. Jodidio <i>et al.</i>	(LBL, NWES, TRIU)
KETOV	86	JETPL 44 146	S.N. Ketov <i>et al.</i>	(KIAE)
KOCH	86	Translated from ZETFP 44 114	H.R. Koch, O.W.B. Schult	(JULI)
KONAKA	86	NC 96A 182	A. Konaka <i>et al.</i>	(KYOT, KEK)
MAIANI	86	PL B175 359	L. Maiani, R. Petronzio, E. Zavattini	(CERN)
PECCEI	86	PL B172 435	R.D. Peccei, T.T. Wu, T. Yanagida	(DESY)
RAFFELT	86	PR D33 897	G.G. Raffelt	(MPIM)
RAFFELT	86B	PL 166B 402	G.G. Raffelt	(MPIM)
SAVAGE	86B	PRL 57 178	M.J. Savage <i>et al.</i>	(CIT)
AMALDI	85	PL 153B 444	U. Amaldi <i>et al.</i>	(CERN)
ANANEV	85	SJNP 41 585	V.D. Ananev <i>et al.</i>	(JINR)
BALTRUSAIT...	85	PRL 55 1842	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BERGSMA	85	PL 157B 458	F. Bergsma <i>et al.</i>	(CHARM Collab.)
KAPLAN	85	NP B260 215	D.B. Kaplan	(HARV)
IWAMOTO	84	PRL 53 1198	N. Iwamoto	(UCSB, WUOL)
YAMAZAKI	84	PRL 52 1089	T. Yamazaki <i>et al.</i>	(INUS, KEK)
ABBOTT	83	PL 120B 133	L.F. Abbott, P. Sikivie	(BRAN, FLOR)
CARBONI	83	PL 123B 349	G. Carbone, W. Dahme	(CERN, MUNI)
CAVAIGNAC	83	PL 121B 193	J.F. Cavaignac <i>et al.</i>	(BNL, LAPP)
DICUS	83	PR D28 1778	D.A. Dicus, V.L. Teplitz	(TEXA, UMD)
DINE	83	PL 120B 137	M. Dine, W. Fischler	(IAS, PERN)
ELLIS	83B	NP B223 252	J. Ellis, K.A. Olive	(CERN)
FAISSNER	83	PR D28 1198	H. Faissner <i>et al.</i>	(AACH)
FAISSNER	83B	PR D28 1787	H. Faissner <i>et al.</i>	(AACH3)
FRANK	83B	PR D28 1790	J.S. Frank <i>et al.</i>	(LANL, YALE, LBL+)
HOFFMAN	83	PR D28 660	C.M. Hoffman <i>et al.</i>	(LANL, ARZS)
PRESKILL	83	PL 120B 127	J. Preskill, M.B. Wise, F. Wilczek	(HARV, UCSBT)
SIKIVIE	83	PRL 51 1415	P. Sikivie	(FLOR)
Also		PRL 52 695 (erratum)	P. Sikivie	(FLOR)
ALEKSEEV	82	JETP 55 591	G.D. Alekseeva <i>et al.</i>	(KIAE)
ALEKSEEV	82B	Translated from ZETFP 82 1007	G.D. Alekseev <i>et al.</i>	(MOSU, JINR)
ASANO	82	JETPL 36 116	Y. Asano <i>et al.</i>	(KEK, TOKY, INUS, OSAK)
Also		Translated from ZETFP 36 94		
ASANO	82	PL 113B 195	Y. Asano <i>et al.</i>	(KEK, TOKY, INUS, OSAK)

BARROSO	82	PL 116B 247	A. Barroso, G.C. Branco	(LISB)
DATAR	82	PL 114B 63	V.M. Datar <i>et al.</i>	(BHAB)
EDWARDS	82	PRL 48 903	C. Edwards <i>et al.</i>	(Crystal Ball Collab.)
FETSCHER	82	JP G8 L147	W. Fetscher	(ETH)
FUKUGITA	82	PRL 48 1522	M. Fukugita, S. Watamura, M. Yoshimura	(KEK)
FUKUGITA	82B	PR D26 1840	M. Fukugita, S. Watamura, M. Yoshimura	(KEK)
LEHMANN	82	PL 115B 270	P. Lehmann <i>et al.</i>	(SACL)
RAFFELT	82	PL 119B 323	G. Raffelt, L. Stodolsky	(MPIM)
ZEHNDRER	82	PL 110B 419	A. Zehnder, K. Gabathuler, J.L. Vuilleumier	(ETH+)
ASANO	81B	PL 107B 159	Y. Asano <i>et al.</i>	(KEK, TOKY, INUS, OSAK)
BARROSO	81	PL 106B 91	A. Barroso, N.C. Mukhopadhyay	(SIN)
FAISSNER	81	ZPHY C10 95	H. Faissner <i>et al.</i>	(AACH3)
FAISSNER	81B	PL 103B 234	H. Faissner <i>et al.</i>	(AACH3)
KIM	81	PL 105B 55	B.R. Kim, C. Stamm	(AACH3)
VUILLEUMIER	81	PL 101B 341	J.L. Vuilleumier <i>et al.</i>	(CIT, MUNI)
ZEHNDRER	81	PL 104B 494	A. Zehnder	(ETH)
FAISSNER	80	PL 96B 201	H. Faissner <i>et al.</i>	(AACH3)
JACQUES	80	PR D21 1206	P.F. Jacques <i>et al.</i>	(RUTG, STEV, COLU)
SOUKAS	80	PRL 44 564	A. Soukas <i>et al.</i>	(BNL, HARV, ORNL, PENN)
BECHIS	79	PRL 42 1511	D.J. Bechis <i>et al.</i>	(UMD, COLU, AFRR)
CALAPRICE	79	PR D20 2708	F.P. Calaprice <i>et al.</i>	(PRIN)
COTEUS	79	PRL 42 1438	P. Coteus <i>et al.</i>	(COLU, ILL, BNL)
DISHAW	79	PL 85B 142	J.P. Dishaw <i>et al.</i>	(SLAC, CIT)
ZHITNITSKII	79	SJNP 29 517	A.R. Zhitnitsky, Y.I. Skovpen	(NOVO)
ALIBRAN	78	PL 74B 134	P. Alibrand <i>et al.</i>	(Gargamelle Collab.)
ASRATYAN	78B	PL 79B 497	A.E. Asratyan <i>et al.</i>	(ITEP, SERP)
BELLOTTI	78	PL 76B 223	E. Bellotti, E. Fiorini, L. Zanotti	(MILA)
BOSETTI	78B	PL 74B 143	P.C. Bosetti <i>et al.</i>	(BEBC Collab.)
DICUS	78C	PR D18 1829	D.A. Dicus <i>et al.</i>	(TEXA, VPI, STAN)
DONNELLY	78	PR D18 1607	T.W. Donnelly <i>et al.</i>	(STAN)
Also		PRL 37 315	F. Reines, H.S. Gurr, H.W. Sobel	(UCI)
HANS	78D	PL 74B 139	H.S. Gurr, F. Reines, H.W. Sobel	(UCI)
MICELMAC...	78	LNC 21 441	G.V. Mitselmakher, B. Pontecorvo	(CDHS Collab.)
MIKAELIAN	78	PR D18 3605	K.O. Mikaelian	(JINR)
SATO	78	PTP 60 1942	K. Sato	(FNAL, NWES)
VYSOTSKII	78	JETPL 27 502	M.I. Vyssotsky <i>et al.</i>	(KYOT)
YANG	78	Translated from ZETFP 27 533	M.I. Vyssotsky <i>et al.</i>	(ASCI)
PECCEI	77	PRL 41 523	T.C. Yang	(MASA)
Also		PR D16 1791	R.D. Peccei, H.R. Quinn	(STAN, SLAC)
REINES	76	PRL 37 315	R.D. Peccei, H.R. Quinn	(STAN, SLAC)
GURR	74	PRL 37 179	F. Reines, H.S. Gurr, H.W. Sobel	(UCI)
ANAND	53	PRSL A22 183	H.S. Gurr, F. Reines, H.W. Sobel	(UCI)
Also			B.M. Anand	

OTHER RELATED PAPERS

SREDNICKI	85	NP B260 689	M. Srednicki	(UCSB)
BARDEEN	78	PL 74B 229	W.A. Bardeen, S.-H.H. Tye	(FNAL)



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## LEPTONS

e

$$J = \frac{1}{2}$$

**e MASS (atomic mass units u)**

The primary determination of an electron's mass comes from measuring the ratio of the mass to that of a nucleus, so that the result is obtained in u (atomic mass units). The conversion factor to MeV is more uncertain than the mass of the electron in u; indeed, the recent improvements in the mass determination are not evident when the result is given in MeV. In this datablock we give the result in u, and in the following datablock in MeV.

VALUE ( $10^{-6}$ u)	DOCUMENT ID	TECN	COMMENT
<b>548.57990970 ± 0.00000016</b>	MOHR	16	RVUE 2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
548.57990946 ± 0.00000022	MOHR	12	RVUE 2010 CODATA value
548.57990943 ± 0.00000023	MOHR	08	RVUE 2006 CODATA value
548.57990945 ± 0.00000024	MOHR	05	RVUE 2002 CODATA value
548.5799092 ± 0.0000004	<sup>1</sup> BEIER	02	CNTR Penning trap
548.5799110 ± 0.0000012	MOHR	99	RVUE 1998 CODATA value
548.5799111 ± 0.0000012	<sup>2</sup> FARNHAM	95	CNTR Penning trap
548.579903 ± 0.000013	COHEN	87	RVUE 1986 CODATA value

<sup>1</sup> BEIER 02 compares Larmor frequency of the electron bound in a  $^{12}\text{C}^{5+}$  ion with the cyclotron frequency of a single trapped  $^{12}\text{C}^{5+}$  ion.

<sup>2</sup> FARNHAM 95 compares cyclotron frequency of trapped electrons with that of a single trapped  $^{12}\text{C}^{6+}$  ion.

**e MASS**

2010 CODATA (MOHR 12) gives the conversion factor from u (atomic mass units, see the above datablock) to MeV as 931.494 061 (21). Earlier values use the then-current conversion factor. The conversion error dominates the uncertainty of the masses given below.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>0.5109989461 ± 0.000000031</b>	MOHR	16	RVUE 2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.510998928 ± 0.000000011	MOHR	12	RVUE 2010 CODATA value
0.510998910 ± 0.000000013	MOHR	08	RVUE 2006 CODATA value
0.510998918 ± 0.000000044	MOHR	05	RVUE 2002 CODATA value
0.510998901 ± 0.000000020	<sup>1,2</sup> BEIER	02	CNTR Penning trap
0.510998902 ± 0.000000021	MOHR	99	RVUE 1998 CODATA value
0.510998903 ± 0.000000020	<sup>1,3</sup> FARNHAM	95	CNTR Penning trap
0.510998895 ± 0.000000024	<sup>1</sup> COHEN	87	RVUE 1986 CODATA value
0.5110034 ± 0.00000014	COHEN	73	RVUE 1973 CODATA value

<sup>1</sup> Converted to MeV using the 1998 CODATA value of the conversion constant, 931.494 013 ± 0.000037 MeV/u.

<sup>2</sup> BEIER 02 compares Larmor frequency of the electron bound in a  $^{12}\text{C}^{5+}$  ion with the cyclotron frequency of a single trapped  $^{12}\text{C}^{5+}$  ion.

<sup>3</sup> FARNHAM 95 compares cyclotron frequency of trapped electrons with that of a single trapped  $^{12}\text{C}^{6+}$  ion.

$$(m_{e^+} - m_{e^-}) / m_{\text{average}}$$

A test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 8 × 10<sup>-9</sup></b>	90	<sup>1</sup> FEE	93	Positronium spectroscopy
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 4 × 10 <sup>-23</sup>	90	<sup>2</sup> DOLGOV	14	From photon mass limit
< 4 × 10 <sup>-8</sup>	90	CHU	84	CNTR Positronium spectroscopy

<sup>1</sup> FEE 93 value is obtained under the assumption that the positronium Rydberg constant is exactly half the hydrogen one.

<sup>2</sup> DOLGOV 14 result is obtained under the assumption that any mass difference between electron and positron would lead to a non-zero photon mass. The PDG 12 limit of  $1 \times 10^{-18}$  eV on the photon mass is in turn used to derive the value quoted here.

$$|q_{e^+} + q_{e^-}|/e$$

A test of *CPT* invariance. See also similar tests involving the proton.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>&lt; 4 × 10<sup>-8</sup></b>	<sup>1</sup> HUGHES	92	RVUE
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 2 × 10 <sup>-18</sup>	<sup>2</sup> SCHAEFER	95	THEO Vacuum polarization
< 1 × 10 <sup>-18</sup>	<sup>3</sup> MUELLER	92	THEO Vacuum polarization

<sup>1</sup> HUGHES 92 uses recent measurements of Rydberg-energy and cyclotron-frequency ratios.

<sup>2</sup> SCHAEFER 95 removes model dependency of MUELLER 92.

<sup>3</sup> MUELLER 92 argues that an inequality of the charge magnitudes would, through higher-order vacuum polarization, contribute to the net charge of atoms.

**e MAGNETIC MOMENT ANOMALY**

$$\mu_e/\mu_B - 1 = (g-2)/2$$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	CHG	COMMENT
<b>1159.65218091 ± 0.00000026</b>	MOHR	16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1159.65218076 ± 0.00000027	MOHR	12	RVUE	2010 CODATA value
1159.65218073 ± 0.00000028	HANNEKE	08	MRS	Single electron
1159.65218111 ± 0.00000074	<sup>1</sup> MOHR	08	RVUE	2006 CODATA value
1159.65218085 ± 0.00000076	<sup>2</sup> ODOM	06	MRS	Single electron
1159.6521859 ± 0.0000038	MOHR	05	RVUE	2002 CODATA value
1159.6521869 ± 0.0000041	MOHR	99	RVUE	1998 CODATA value
1159.652193 ± 0.000010	COHEN	87	RVUE	1986 CODATA value
1159.6521884 ± 0.0000043	VANDYCK	87	MRS	Single electron
1159.6521879 ± 0.0000043	VANDYCK	87	MRS	Single positron

<sup>1</sup> MOHR 08 average is dominated by ODOM 06.

<sup>2</sup> Superseded by HANNEKE 08 per private communication with Gerald Gabrielse.

$$(g_{e^+} - g_{e^-}) / g_{\text{average}}$$

A test of *CPT* invariance.

VALUE (units $10^{-12}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>- 0.5 ± 2.1</b>	90	<sup>1</sup> VANDYCK	87	MRS Penning trap
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 12	95	<sup>2</sup> VASSERMAN	87	CNTR Assumes $m_{e^+} = m_{e^-}$
22 ± 64		SCHWINBERG	81	MRS Penning trap

<sup>1</sup> VANDYCK 87 measured  $(g_-/g_+) - 1$  and we converted it.

<sup>2</sup> VASSERMAN 87 measured  $(g_+ - g_-)/(g-2)$ . We multiplied by  $(g-2)/g = 1.2 \times 10^{-3}$ .

**e ELECTRIC DIPOLE MOMENT (d)**

A nonzero value is forbidden by both *T* invariance and *P* invariance.

VALUE ( $10^{-28}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.87</b>	90	<sup>1</sup> BARON	14	CNTR ThO molecules
• • • We do not use the following data for averages, fits, limits, etc. • • •				
- 5570 ± 7980 ± 120		KIM	15	CNTR Gd <sub>3</sub> Ga <sub>5</sub> O <sub>12</sub> molecules
< 6050	90	<sup>2</sup> ECKEL	12	CNTR Eu <sub>0.5</sub> Ba <sub>0.5</sub> TiO <sub>3</sub> molecules
< 10.5	90	<sup>3</sup> HUDSON	11	NMR YbF molecules
6.9 ± 7.4		REGAN	02	MRS 205Ti beams
18 ± 12 ± 10		<sup>4</sup> COMMINS	94	MRS 205Ti beams
- 27 ± 83		<sup>4</sup> ABDULLAH	90	MRS 205Ti beams
- 1400 ± 2400		CHO	89	NMR TlF molecules
- 150 ± 550 ± 150		MURTHY	89	Cs, no B field
- 5000 ± 11000		LAMOREAUX	87	NMR <sup>199</sup> Hg
19000 ± 34000	90	SANDARS	75	MRS Thallium
7000 ± 22000	90	PLAYER	70	MRS Xenon
< 30000	90	WEISSKOPF	68	MRS Cesium

<sup>1</sup> BARON 14 gives a measurement corresponding to this limit as  $(-0.21 \pm 0.37 \pm 0.25) \times 10^{-28}$  ecm.

<sup>2</sup> ECKEL 12 gives a measurement corresponding to this limit as  $(-1.07 \pm 3.06 \pm 1.74) \times 10^{-25}$  ecm.

<sup>3</sup> HUDSON 11 gives a measurement corresponding to this limit as  $(-2.4 \pm 5.7 \pm 1.5) \times 10^{-28}$  ecm.

<sup>4</sup> ABDULLAH 90, COMMINS 94, and REGAN 02 use the relativistic enhancement of a valence electron's electric dipole moment in a high-Z atom.

**e<sup>-</sup> MEAN LIFE / BRANCHING FRACTION**

A test of charge conservation. See the "Note on Testing Charge Conservation and the Pauli Exclusion Principle" following this section in our 1992 edition (Physical Review **D45** 51 (1992), p. VI.10).

Most of these experiments are one of three kinds: Attempts to observe (a) the 255.5 keV gamma ray produced in  $e^- \rightarrow \nu_e \gamma$ , (b) the (K) shell x-ray produced when an electron decays without additional energy deposit, e.g.,  $e^- \rightarrow \nu_e \bar{\nu}_e \nu_e$  ("disappearance" experiments), and (c) nuclear de-excitation gamma rays after the electron disappears from an atomic shell and the nucleus is left in an excited state. The last can include both weak boson and photon mediating processes. We use the best  $e^- \rightarrow \nu_e \gamma$  limit for the Summary Tables.

Note that we use the mean life rather than the half life, which is often reported.

**e → ν<sub>e</sub> γ and astrophysical limits**

VALUE (yr)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt; 6.6 × 10<sup>28</sup></b>	90	AGOSTINI	15B	BORX $e^- \rightarrow \nu \gamma$

# Lepton Particle Listings

*e*, *μ*

• • • We do not use the following data for averages, fits, limits, etc. • • •

>1.22 × 10 <sup>26</sup>	68	<sup>1</sup> KLAPDOR-K...	07	CNTR	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i>
>4.6 × 10 <sup>26</sup>	90	BACK	02	BORX	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i>
>3.4 × 10 <sup>26</sup>	68	BELLI	00B	DAMA	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i> , liquid Xe
>3.7 × 10 <sup>25</sup>	68	AHARONOV	95B	CNTR	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i>
>2.35 × 10 <sup>25</sup>	68	BALYSH	93	CNTR	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i> , <sup>76</sup> Ge detector
>1.5 × 10 <sup>25</sup>	68	AVIGNONE	86	CNTR	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i>
>1 × 10 <sup>39</sup>		<sup>2</sup> ORITO	85	ASTR	Astrophysical argument
>3 × 10 <sup>23</sup>	68	BELLOTTI	83B	CNTR	<i>e</i> <sup>−</sup> → <i>ν</i> <i>γ</i>

<sup>1</sup> The authors of A. Derbin et al, arXiv:0704.2047v1 argue that this limit is overestimated by at least a factor of 5.

<sup>2</sup> ORITO 85 assumes that electromagnetic forces extend out to large enough distances and that the age of our galaxy is 10<sup>10</sup> years.

## Disappearance and nuclear-de-excitation experiments

VALUE (yr)	CL%	DOCUMENT ID	TECN	COMMENT
>6.4 × 10 <sup>24</sup>	68	<sup>1</sup> BELLI	99B	DAMA De-excitation of <sup>129</sup> Xe
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>4.2 × 10 <sup>24</sup>	68	BELLI	99	DAMA Iodine L-shell disappearance
>2.4 × 10 <sup>23</sup>	90	<sup>2</sup> BELLI	99D	DAMA De-excitation of <sup>127</sup> I (in NaI)
>4.3 × 10 <sup>23</sup>	68	AHARONOV	95B	CNTR Ge K-shell disappearance
>2.7 × 10 <sup>23</sup>	68	REUSSER	91	CNTR Ge K-shell disappearance
>2 × 10 <sup>22</sup>	68	BELLOTTI	83B	CNTR Ge K-shell disappearance

<sup>1</sup> BELLI 99B limit on charge nonconserving *e*<sup>−</sup> capture involving excitation of the 236.1 keV nuclear state of <sup>129</sup>Xe; the 90% CL limit is 3.7 × 10<sup>24</sup> yr. Less stringent limits for other states are also given.

<sup>2</sup> BELLI 99D limit on charge nonconserving *e*<sup>−</sup> capture involving excitation of the 57.6 keV nuclear state of <sup>127</sup>I. Less stringent limits for the other states and for the state of <sup>23</sup>Na are also given.

## LIMITS ON LEPTON-FLAVOR VIOLATION IN PRODUCTION

Forbidden by lepton family number conservation.

This section was added for the 2008 edition of this *Review* and is not complete. For a list of further measurements see references in the papers listed below.

### *σ*(*e*<sup>+</sup> *e*<sup>−</sup> → *e*<sup>±</sup> *τ*<sup>∓</sup>) / *σ*(*e*<sup>+</sup> *e*<sup>−</sup> → *μ*<sup>±</sup> *μ*<sup>−</sup>)

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<8.9 × 10 <sup>−6</sup>	95	AUBERT	07P	BABR <i>e</i> <sup>+</sup> <i>e</i> <sup>−</sup> at <i>E</i> <sub>cm</sub> = 10.58 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.8 × 10 <sup>−3</sup>	95	GOMEZ-CAD...	91	MRK2 <i>e</i> <sup>+</sup> <i>e</i> <sup>−</sup> at <i>E</i> <sub>cm</sub> = 29 GeV

### *σ*(*e*<sup>+</sup> *e*<sup>−</sup> → *μ*<sup>±</sup> *τ*<sup>∓</sup>) / *σ*(*e*<sup>+</sup> *e*<sup>−</sup> → *μ*<sup>±</sup> *μ*<sup>−</sup>)

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<4.0 × 10 <sup>−6</sup>	95	AUBERT	07P	BABR <i>e</i> <sup>+</sup> <i>e</i> <sup>−</sup> at <i>E</i> <sub>cm</sub> = 10.58 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<6.1 × 10 <sup>−3</sup>	95	GOMEZ-CAD...	91	MRK2 <i>e</i> <sup>+</sup> <i>e</i> <sup>−</sup> at <i>E</i> <sub>cm</sub> = 29 GeV

## e REFERENCES

MOHR	16	arXiv:1507.07956	P.J. Mohr, D.B. Newell, B.N. Taylor	(NIST)
Accepted for publication in RMP				
AGOSTINI	15B	PRL 115 231802	M. Agostini et al.	(BOREXINO Collab.)
KIM	15	PR D91 102004	Y.J. Kim et al.	(IND, YALE, LANL)
BARON	14	SCIENCE 343 269	J. Baron et al.	(ACME Collab.)
DOLGOV	14	PL B732 244	A.D. Dolgov, V.A. Novikov	
ECKEL	12	PRL 109 193003	S. Eckel, A.O. Sushkov, S.K. Lamoreaux	(YALE)
MOHR	12	RMP 84 1527	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
PDG	12	PR D86 010001	J. Beringer et al.	(PDG Collab.)
HUDSON	11	NAT 473 493	J.J. Hadson et al.	(LOIC)
HANNEKE	08	PRL 100 120801	D. Hanneke, S. Fogwell, G. Gabrielse	(HARV)
MOHR	08	RMP 80 633	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
AUBERT	07P	PR D75 031103	B. Aubert et al.	(BABAR Collab.)
KLAPDOR-K...	07	PL B644 109	H.V. Klapdor-Kleingrothaus, I.V. Krivosheina, I.V. Titkova	(HARV)
ODOM	06	PRL 97 030801	B. Odom et al.	(NIST)
MOHR	05	RMP 77 1	P.J. Mohr, B.N. Taylor	(NIST)
BACK	02	PL B525 29	H.O. Back et al.	(BOREXINO/SASSO Collab.)
BEIER	02	PRL 88 011603	T. Beier et al.	
REGAN	02	PRL 88 071805	B.C. Regan et al.	
BELLI	00B	PR D61 117301	P. Belli et al.	(DAMA Collab.)
BELLI	99	PL B460 236	P. Belli et al.	(DAMA Collab.)
BELLI	99B	PL B465 315	P. Belli et al.	(DAMA Collab.)
BELLI	99D	PR C60 065501	P. Belli et al.	(DAMA Collab.)
MOHR	99	JPCRD 28 1713	P.J. Mohr, B.N. Taylor	(NIST)
Also		RMP 72 351	P.J. Mohr, B.N. Taylor	(NIST)
AHARONOV	95B	PR D52 3785	Y. Aharonov et al.	(SCUC, PNL, ZARA+)
Also		PL B353 168	Y. Aharonov et al.	(SCUC, PNL, ZARA+)
FARNHAM	95	PR 75 3598	D.L. Farnham, R.S. van Dyck, P.B. Schwinberg	(WASH)
SCHAEFER	95	PR A51 838	A. Schaefer, J. Reinhardt	(FRAN)
COMMINS	94	PR A50 2960	E.D. Commins et al.	
BALYSH	93	PL B298 278	A. Balysh et al.	(KIAE, MPHI, SASSO)
FEE	93	PR A48 192	M.S. Fee et al.	
HUGHES	92	PRL 69 578	R.J. Hughes, B.I. Deutch	(LANL, AARH)
MUELLER	92	PRL 69 3432	B. Muller, M.H. Thoma	(DUKE)
PDG	92	PR D45 51	K. Hikasa et al.	(KEK, LBL, BOST+)
GOMEZ-CAD...	91	PRL 66 1007	J.J. Gomez-Cadenas et al.	(SLAC MARK-2 Collab.)
REUSSER	91	PL B255 143	D. Reusser et al.	(NEUC, CIT, PSI)
ABDULLAH	90	PRL 65 2347	K. Abdullah et al.	(LBL, UCB)
CHO	89	PRL 63 3559	D. Cho, K. Sangster, E.A. Hinds	(YALE)
MURTHY	89	PR 63 965	S.A. Murthy et al.	(AMHT)
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RISC, NBS)
LAMOREAUX	87	PRL 59 2275	S.K. Lamoreaux et al.	(WASH)
VANDYCK	87	PRL 59 26	R.S. van Dyck, P.B. Schwinberg, H.G. Dehmelt	(WASH)
VASSERMAN	87	PL B198 302	I.B. Vasserman et al.	(NOVO)
Also		PL B187 172	I.B. Vasserman et al.	(NOVO)

AVIGNONE	86	PR D34 97	F.T. Avignone et al.	(PNL, SCUC)
ORITO	85	PRL 54 2457	S. Orto, M. Yoshimura	(TOKY, KEK)
CHU	84	PRL 52 1689	S. Chu, A.P. Mills, J.L. Hall	(BELL, NBS, COLO)
BELLOTTI	83B	PL 124B 435	E. Bellotti et al.	(MILA)
SCHWINBERG	81	PRL 47 1679	P.B. Schwinberg, R.S. van Dyck, H.G. Dehmelt	(WASH)
SANDARS	75	PR A11 473	P.G.H. Sandars, D.M. Sternheimer	(OXF, BNL)
COHEN	73	JPCRD 2 664	E.R. Cohen, B.N. Taylor	(RISC, NBS)
PLAYER	70	JP B3 1620	M.A. Player, P.G.H. Sandars	(OXF)
WEISSKOPF	68	PRL 21 1645	M.C. Weisskopf et al.	(BRAN)



$$J = \frac{1}{2}$$

## μ MASS (atomic mass units u)

The muon's mass is obtained from the muon-electron mass ratio as determined from the measurement of Zeeman transition frequencies in muonium (*μ*<sup>+</sup> *e*<sup>−</sup> atom). Since the electron's mass is most accurately known in *u*, the muon's mass is also most accurately known in *u*. The conversion factor to MeV has approximately the same relative uncertainty as the mass of the muon in *u*. In this datablock we give the result in *u*, and in the following datablock in MeV.

VALUE (u)	DOCUMENT ID	TECN	COMMENT
<b>0.1134289257 ± 0.0000000025</b>	MOHR	16	RVUE 2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1134289267 ± 0.0000000029	MOHR	12	RVUE 2010 CODATA value
0.1134289256 ± 0.0000000029	MOHR	08	RVUE 2006 CODATA value
0.1134289264 ± 0.0000000030	MOHR	05	RVUE 2002 CODATA value
0.1134289168 ± 0.0000000034	<sup>1</sup> MOHR	99	RVUE 1998 CODATA value
0.113428913 ± 0.0000000017	<sup>2</sup> COHEN	87	RVUE 1986 CODATA value

<sup>1</sup> MOHR 99 make use of other 1998 CODATA entries below.

<sup>2</sup> COHEN 87 make use of other 1986 CODATA entries below.

## μ MASS

2010 CODATA (MOHR 12) gives the conversion factor from *u* (atomic mass units, see the above datablock) to MeV as 931.494 061 (21). Earlier values use the then-current conversion factor. The conversion error contributes significantly to the uncertainty of the masses given below.

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>105.6583745 ± 0.00000024</b>	MOHR	16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •				
105.6583715 ± 0.00000035	MOHR	12	RVUE	2010 CODATA value
105.6583668 ± 0.00000038	MOHR	08	RVUE	2006 CODATA value
105.6583692 ± 0.00000094	MOHR	05	RVUE	2002 CODATA value
105.6583568 ± 0.00000052	MOHR	99	RVUE	1998 CODATA value
105.658353 ± 0.0000016	<sup>1</sup> COHEN	87	RVUE	1986 CODATA value
105.658386 ± 0.0000044	<sup>2</sup> MARIAM	82	CNTR +	
105.65836 ± 0.000026	<sup>3</sup> CROWE	72	CNTR	
105.65865 ± 0.00044	<sup>4</sup> CRANE	71	CNTR	

<sup>1</sup> Converted to MeV using the 1998 CODATA value of the conversion constant, 931.494013 ± 0.000037 MeV/*u*.

<sup>2</sup> MARIAM 82 give *m*<sub>μ</sub>/*m*<sub>*e*</sub> = 206.768259(62).

<sup>3</sup> CROWE 72 give *m*<sub>μ</sub>/*m*<sub>*e*</sub> = 206.7682(5).

<sup>4</sup> CRANE 71 give *m*<sub>μ</sub>/*m*<sub>*e*</sub> = 206.76878(85).

## μ MEAN LIFE τ

Measurements with an error > 0.001 × 10<sup>−6</sup> s have been omitted.

VALUE (10 <sup>−6</sup> s)	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.1969811 ± 0.0000022 OUR AVERAGE</b>				
2.1969803 ± 0.00000021 ± 0.0000007 <sup>1</sup>	TISHCHENKO	13	CNTR	+ Surface <i>μ</i> <sup>+</sup> at PSI
2.197083 ± 0.000032 ± 0.000015	BARCZYK	08	CNTR	+ Muons from <i>π</i> <sup>+</sup> decay at rest
2.197013 ± 0.000021 ± 0.000011	CHITWOOD	07	CNTR	+ Surface <i>μ</i> <sup>+</sup> at PSI
2.197078 ± 0.000073	BARDIN	84	CNTR	+
2.197025 ± 0.000155	BARDIN	84	CNTR	−
2.19695 ± 0.00006	GIOVANETTI	84	CNTR	+
2.19711 ± 0.00008	BALANDIN	74	CNTR	+
2.1973 ± 0.0003	DUCLIOS	73	CNTR	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.1969803 ± 0.0000022	WEBBER	11	CNTR	+ Surface <i>μ</i> <sup>+</sup> at PSI
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<sup>1</sup> TISHCHENKO 13 uses 1.6 × 10<sup>12</sup> *μ*<sup>+</sup> events and supersedes WEBBER 11.

## τ<sub>μ<sup>+</sup></sub>/τ<sub>μ<sup>−</sup></sub> MEAN LIFE RATIO

A test of *CPT* invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.000024 ± 0.000078</b>	BARDIN	84	CNTR
• • • We do not use the following data for averages, fits, limits, etc. • • •			

1.0008	$\pm 0.0010$	BAILEY	79	CNTR	Storage ring
1.000	$\pm 0.001$	MEYER	63	CNTR	Mean life $\mu^+ / \mu^-$

$$(\tau_{\mu^+} - \tau_{\mu^-}) / \tau_{\text{average}}$$

A test of *CPT* invariance. Calculated from the mean-life ratio, above.

VALUE	DOCUMENT ID
$(2 \pm 8) \times 10^{-5}$	OUR EVALUATION

### $\mu/p$ MAGNETIC MOMENT RATIO

This ratio is used to obtain a precise value of the muon mass and to reduce experimental muon Larmor frequency measurements to the muon magnetic moment anomaly. Measurements with an error  $> 0.00001$  have been omitted. By convention, the minus sign on this ratio is omitted. CODATA values were fitted using their selection of data, plus other data from multiparameter fits.

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>3.183345142 <math>\pm 0.000000071</math></b>	MOHR	16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.183345107 $\pm 0.000000084$	MOHR	12	RVUE	2010 CODATA value
3.183345137 $\pm 0.000000085$	MOHR	08	RVUE	2006 CODATA value
3.183345118 $\pm 0.000000089$	MOHR	05	RVUE	2002 CODATA value
3.18334513 $\pm 0.000000039$	LIU	99	CNTR +	HFS in muonium
3.18334539 $\pm 0.000000010$	MOHR	99	RVUE	1998 CODATA value
3.18334547 $\pm 0.000000047$	COHEN	87	RVUE	1986 CODATA value
3.1833441 $\pm 0.00000017$	KLEMPPT	82	CNTR +	Precession strob
3.1833461 $\pm 0.00000011$	MARIAM	82	CNTR +	HFS splitting
3.1833448 $\pm 0.00000029$	CAMANI	78	CNTR +	See KLEMPPT 82
3.1833403 $\pm 0.00000044$	CASPERSON	77	CNTR +	HFS splitting
3.1833402 $\pm 0.00000072$	COHEN	73	RVUE	1973 CODATA value
3.1833467 $\pm 0.00000082$	CROWE	72	CNTR +	Precession phase

### THE MUON ANOMALOUS MAGNETIC MOMENT

Updated August 2013 by A. Hoecker (CERN), and W.J. Marciano (BNL).

The Dirac equation predicts a muon magnetic moment,  $\vec{M} = g_\mu \frac{e}{2m_\mu} \vec{S}$ , with gyromagnetic ratio  $g_\mu = 2$ . Quantum loop effects lead to a small calculable deviation from  $g_\mu = 2$ , parameterized by the anomalous magnetic moment

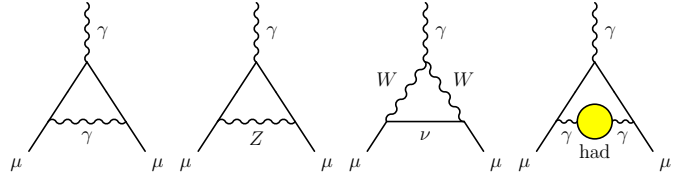
$$a_\mu \equiv \frac{g_\mu - 2}{2}. \quad (1)$$

That quantity can be accurately measured and, within the Standard Model (SM) framework, precisely predicted. Hence, comparison of experiment and theory tests the SM at its quantum loop level. A deviation in  $a_\mu^{\text{exp}}$  from the SM expectation would signal effects of new physics, with current sensitivity reaching up to mass scales of  $\mathcal{O}(\text{TeV})$  [1,2]. For recent and very thorough muon  $g - 2$  reviews, see Refs. [3–5].

The E821 experiment at Brookhaven National Lab (BNL) studied the precession of  $\mu^+$  and  $\mu^-$  in a constant external magnetic field as they circulated in a confining storage ring. It found [7]<sup>1</sup>

$$\begin{aligned} a_{\mu^+}^{\text{exp}} &= 11\,659\,204(6)(5) \times 10^{-10}, \\ a_{\mu^-}^{\text{exp}} &= 11\,659\,215(8)(3) \times 10^{-10}, \end{aligned} \quad (2)$$

<sup>1</sup> The original results reported by the experiment have been updated in Eq. (2) and Eq. (3) to the newest value for the absolute muon-to-proton magnetic ratio  $\lambda = 3.183\,345\,107(84)$  [6]. The change induced in  $a_\mu^{\text{exp}}$  with respect to the value of  $\lambda = 3.183\,345\,39(10)$  used in Ref. 7 amounts to  $+1.12 \times 10^{-10}$ .



where the first errors are statistical and the second systematic. Assuming *CPT* invariance and taking into account correlations between systematic uncertainties, one finds for their average [6,7]

$$a_\mu^{\text{exp}} = 11\,659\,209.1(5.4)(3.3) \times 10^{-10}. \quad (3)$$

These results represent about a factor of 14 improvement over the classic CERN experiments of the 1970's [8]. Improvement of the measurement in Eq. (3) by a factor of four by moving the E821 storage ring to Fermilab, and utilizing a cleaner and more intense muon beam is in progress. An even more ambitious precision goal is set by an experiment based on a beam of ultra-cold muons proposed at the Japan Proton Accelerator Research Complex.

The SM prediction for  $a_\mu^{\text{SM}}$  is generally divided into three parts (see Fig. 1 for representative Feynman diagrams)

$$a_\mu^{\text{SM}} = a_\mu^{\text{QED}} + a_\mu^{\text{EW}} + a_\mu^{\text{Had}}. \quad (4)$$

The QED part includes all photonic and leptonic ( $e, \mu, \tau$ ) loops starting with the classic  $\alpha/2\pi$  Schwinger contribution. It has been computed through 5 loops [9]

$$\begin{aligned} a_\mu^{\text{QED}} &= \frac{\alpha}{2\pi} + 0.765\,857\,425(17) \left(\frac{\alpha}{\pi}\right)^2 + 24.050\,509\,96(32) \left(\frac{\alpha}{\pi}\right)^3 \\ &\quad + 130.879\,6(6.3) \left(\frac{\alpha}{\pi}\right)^4 + 753.3(1.0) \left(\frac{\alpha}{\pi}\right)^5 + \dots \end{aligned} \quad (5)$$

with a few significant changes in the coefficients since our previous update of this review in 2011. Employing<sup>2</sup>  $\alpha^{-1} = 137.035\,999\,049(90)$ , obtained [6] from the precise measurements of  $h/m_{\text{Rb}}$  [11], the Rydberg constant and  $m_{\text{Rb}}/m_e$  [6], leads to [9]

$$a_\mu^{\text{QED}} = 116\,584\,718.95(0.08) \times 10^{-11}, \quad (6)$$

where the small error results mainly from the uncertainty in  $\alpha$ .

Loop contributions involving heavy  $W^\pm, Z$  or Higgs particles are collectively labeled as  $a_\mu^{\text{EW}}$ . They are suppressed by at least a factor of  $\frac{\alpha}{\pi} \frac{m_\mu^2}{m_W^2} \simeq 4 \times 10^{-9}$ . At 1-loop order [12]

$$\begin{aligned} a_\mu^{\text{EW}}[1\text{-loop}] &= \frac{G_\mu m_\mu^2}{8\sqrt{2}\pi^2} \left[ \frac{5}{3} + \frac{1}{3} (1 - 4\sin^2\theta_W)^2 \right. \\ &\quad \left. + \mathcal{O}\left(\frac{m_\mu^2}{M_W^2}\right) + \mathcal{O}\left(\frac{m_\mu^2}{M_H^2}\right) \right], \\ &= 194.8 \times 10^{-11}, \end{aligned} \quad (7)$$

<sup>2</sup> In the previous versions of this review we used the precise  $\alpha$  value determined from the electron  $a_e$  measurement [9,10]. With the new measurement [11] of the recoil velocity of Rubidium,  $h/m_{\text{Rb}}$ , an  $a_e$ -independent determination of  $\alpha$  with sufficient precision is available and preferred.

# Lepton Particle Listings

$\mu$

for  $\sin^2\theta_W \equiv 1 - M_W^2/M_Z^2 \simeq 0.223$ , and where  $G_\mu \simeq 1.166 \times 10^{-5} \text{ GeV}^{-2}$  is the Fermi coupling constant. Two-loop corrections are relatively large and negative [13]. For a Higgs boson mass of  $\simeq 126 \text{ GeV}$  [13]

$$a_\mu^{\text{EW}}[\text{2-loop}] = -41.2(1.0) \times 10^{-11}, \quad (8)$$

where the uncertainty stems from quark triangle loops. The 3-loop leading logarithms are negligible [13,14],  $\mathcal{O}(10^{-12})$ , implying in total

$$a_\mu^{\text{EW}} = 153.6(1.0) \times 10^{-11}. \quad (9)$$

Hadronic (quark and gluon) loop contributions to  $a_\mu^{\text{SM}}$  give rise to its main theoretical uncertainties. At present, those effects are not calculable from first principles, but such an approach, at least partially, may become possible as lattice QCD matures. Instead, one currently relies on a dispersion relation approach to evaluate the lowest-order (*i.e.*,  $\mathcal{O}(\alpha^2)$ ) hadronic vacuum polarization contribution  $a_\mu^{\text{Had}}[\text{LO}]$  from corresponding cross section measurements [15]

$$a_\mu^{\text{Had}}[\text{LO}] = \frac{1}{3} \left( \frac{\alpha}{\pi} \right)^2 \int_{m_\pi^2}^{\infty} ds \frac{K(s)}{s} R^{(0)}(s), \quad (10)$$

where  $K(s)$  is a QED kernel function [16], and where  $R^{(0)}(s)$  denotes the ratio of the bare<sup>3</sup> cross section for  $e^+e^-$  annihilation into hadrons to the pointlike muon-pair cross section at center-of-mass energy  $\sqrt{s}$ . The function  $K(s) \sim 1/s$  in Eq. (10) gives a strong weight to the low-energy part of the integral. Hence,  $a_\mu^{\text{Had}}[\text{LO}]$  is dominated by the  $\rho(770)$  resonance.

Currently, the available  $\sigma(e^+e^- \rightarrow \text{hadrons})$  data give a leading-order hadronic vacuum polarization (representative) contribution of [17]

$$a_\mu^{\text{Had}}[\text{LO}] = 6923(42)(3) \times 10^{-11}, \quad (11)$$

where the first error is experimental (dominated by systematic uncertainties), and the second due to perturbative QCD, which is used at intermediate and large energies to predict the contribution from the quark-antiquark continuum. New multi-hadron data from the BABAR experiment have increased the constraints on unmeasured exclusive final states and led to a small reduction in the hadronic contribution compared to the 2009 PDG value.

Alternatively, one can use precise vector spectral functions from  $\tau \rightarrow \nu_\tau + \text{hadrons}$  decays [18] that can be related to isovector  $e^+e^- \rightarrow \text{hadrons}$  cross sections by isospin symmetry. Replacing  $e^+e^-$  data in the two-pion and four-pion channels by the corresponding isospin-transformed  $\tau$  data, and applying

<sup>3</sup> The bare cross section is defined as the measured cross section corrected for initial-state radiation, electron-vertex loop contributions and vacuum-polarization effects in the photon propagator. However, QED effects in the hadron vertex and final state, as photon radiation, are included.

isospin-violating corrections (from QED and  $m_d - m_u \neq 0$ ), one finds [17]

$$a_\mu^{\text{Had}}[\text{LO}] = 7015(42)(19)(3) \times 10^{-11} (\tau), \quad (12)$$

where the first error is experimental, the second estimates the uncertainty in the isospin-breaking corrections applied to the  $\tau$  data, and the third error is due to perturbative QCD. The current discrepancy between the  $e^+e^-$  and  $\tau$ -based determinations of  $a_\mu^{\text{Had}}[\text{LO}]$  has been reduced to  $1.8\sigma$  with respect to earlier evaluations. New  $e^+e^-$  and  $\tau$  data from the  $B$ -factory experiments BABAR and Belle have increased the experimental information. Reevaluated isospin-breaking corrections have also contributed to this improvement [19]. BABAR reported good agreement with the  $\tau$  data in the most important two-pion channel [20]. The remaining discrepancy with the older  $e^+e^-$  and  $\tau$  datasets may be indicative of problems with one or both data sets. It may also suggest the need for additional isospin-violating corrections to the  $\tau$  data. Several evaluations of  $a_\mu^{\text{Had}}[\text{LO}]$  have been published leading to similar results (see Fig. 2). The low-energy contribution to  $a_\mu^{\text{Had}}[\text{LO}]$  has also been evaluated with the use of additional theory or model constraints in Refs. [22] and [23], respectively.

Higher order,  $\mathcal{O}(\alpha^3)$ , hadronic contributions are obtained from dispersion relations using the same  $e^+e^- \rightarrow \text{hadrons}$  data [18,21,24], giving  $a_\mu^{\text{Had,Disp}}[\text{NLO}] = (-98.4 \pm 0.6) \times 10^{-11}$ , along with model-dependent estimates of the hadronic light-by-light scattering contribution,  $a_\mu^{\text{Had,LBL}}[\text{NLO}]$ , motivated by large- $N_C$  QCD [25–31].<sup>4</sup> Following [29], one finds for the sum of the two terms

$$a_\mu^{\text{Had}}[\text{NLO}] = 7(26) \times 10^{-11}, \quad (13)$$

where the error is dominated by hadronic light-by-light uncertainties.

Adding Eqs. (6), (9), (11) and (13) gives the representative  $e^+e^-$  data based SM prediction

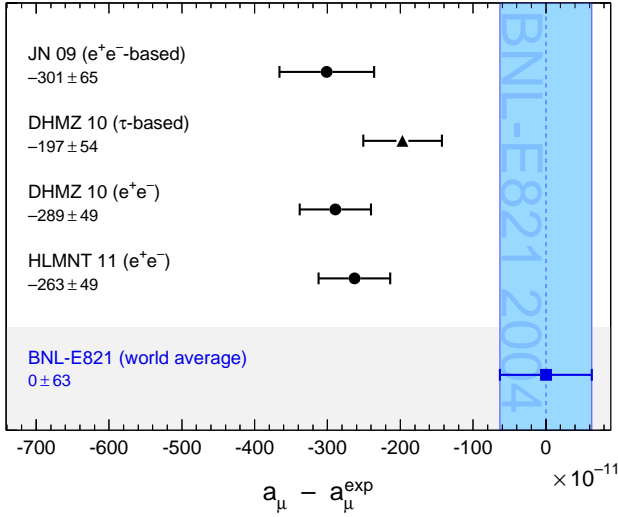
$$a_\mu^{\text{SM}} = 116591803(1)(42)(26) \times 10^{-11}, \quad (14)$$

where the errors are due to the electroweak, lowest-order hadronic, and higher-order hadronic contributions, respectively. The difference between experiment and theory

$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = 288(63)(49) \times 10^{-11}, \quad (15)$$

(with all errors combined in quadrature) represents an interesting but not yet conclusive discrepancy of 3.6 times the estimated  $1\sigma$  error. All the recent estimates for the hadronic contribution compiled in Fig. 2 exhibit similar discrepancies. Switching to  $\tau$  data reduces the discrepancy to  $2.4\sigma$ , assuming the isospin-violating corrections are under control within the

<sup>4</sup> Some representative recent estimates of the hadronic light-by-light scattering contribution,  $a_\mu^{\text{Had,LBL}}[\text{NLO}]$ , that followed after the sign correction of [27], are:  $105(26) \times 10^{-11}$  [29],  $110(40) \times 10^{-11}$  [25],  $136(25) \times 10^{-11}$  [26].



**Figure 2:** Compilation of recent published results for  $a_\mu$  (in units of  $10^{-11}$ ), subtracted by the central value of the experimental average (3). The shaded band indicates the size of the experimental uncertainty. The SM predictions are taken from: JN [4], DHMZ [17], HMNT [21]. Note that the quoted errors in the figure do not include the uncertainty on the subtracted experimental value. To obtain for each theory calculation a result equivalent to Eq. (15), the errors from theory and experiment must be added in quadrature.

estimated uncertainties (see Ref. 32 for an analysis leading to a different conclusion).

An alternate interpretation is that  $\Delta a_\mu$  may be a new physics signal with supersymmetric particle loops as the leading candidate explanation. Such a scenario is quite natural, since generically, supersymmetric models predict [1] an additional contribution to  $a_\mu^{\text{SM}}$

$$a_\mu^{\text{SUSY}} \simeq \text{sign}(\mu) \cdot 130 \times 10^{-11} \cdot \left( \frac{100 \text{ GeV}}{m_{\text{SUSY}}} \right)^2 \tan\beta, \quad (16)$$

where  $m_{\text{SUSY}}$  is a representative supersymmetric mass scale,  $\tan\beta \simeq 3\text{--}40$  a potential enhancement factor, and  $\text{sign}(\mu) = \pm 1$ . Supersymmetric particles in the mass range 100–500 GeV could be the source of the deviation  $\Delta a_\mu$ . If so, those particles should be directly observed at the Large Hadron Collider at CERN.

New physics effects [1] other than supersymmetry could also explain a non-vanishing  $\Delta a_\mu$ . A recent popular scenario involves the “dark photon”, a relatively light hypothetical vector boson from the dark matter sector that couples to our world of particle physics through mixing with the ordinary photon [33–35]. As a result, it couples to ordinary charged particles with strength  $\varepsilon \cdot e$  and gives rise to an additional muon anomalous magnetic moment contribution

$$a_\mu^{\text{dark photon}} = \frac{\alpha}{2\pi} \varepsilon^2 F(m_V/m_\mu), \quad (17)$$

where  $F(x) = \int_0^1 2z(1-z)^2 / [(1-z)^2 + x^2 z] dz$ . For values of  $\varepsilon \sim 1\text{--}2 \cdot 10^{-3}$  and  $m_V \sim 10\text{--}100 \text{ MeV}$ , the dark photon, which

was originally motivated by cosmology, can provide a viable solution to the muon  $g - 2$  discrepancy. Searches for the dark photon in that mass range are currently underway at Jefferson Lab, USA, and MAMI in Mainz, Germany.

## References

1. A. Czarnecki and W.J. Marciano, Phys. Rev. **D64**, 013014 (2001).
2. M. Davier and W.J. Marciano, Ann. Rev. Nucl. and Part. Sci. **54**, 115 (2004).
3. J. Miller, E. de Rafael, and B. Lee Roberts, Rept. Prog. Phys. **70**, 795 (2007).
4. F. Jegerlehner and A. Nyffeler, Phys. Reports **477**, 1 (2009).
5. J.P. Miller *et al.*, Ann. Rev. Nucl. and Part. Sci. **62**, 237 (2012).
6. P.J. Mohr, B.N. Taylor, and D.B. Newell, CODATA Group, Rev. Mod. Phys. **84**, 1527 (2012).
7. G.W. Bennett *et al.*, Phys. Rev. Lett. **89**, 101804 (2002); Erratum *ibid.* Phys. Rev. Lett. **89**, 129903 (2002); G.W. Bennett *et al.*, Phys. Rev. Lett. **92**, 161802 (2004); G.W. Bennett *et al.*, Phys. Rev. Lett. **D73**, 072003 (2006).
8. J. Bailey *et al.*, Nucl. Phys. **B150**, 1 (1979).
9. T. Aoyama *et al.*, Phys. Rev. Lett. **109**, 111808 (2012); T. Aoyama *et al.*, Phys. Rev. Lett. **109**, 111807 (2012); T. Kinoshita and M. Nio, Phys. Rev. **D73**, 013003 (2006); T. Aoyama *et al.*, Phys. Rev. Lett. **99**, 110406 (2007); T. Kinoshita and M. Nio, Phys. Rev. **D70**, 113001 (2004); T. Kinoshita, Nucl. Phys. **B144**, 206 (2005)(Proc. Supp.); T. Kinoshita and M. Nio, Phys. Rev. **D73**, 053007 (2006); A.L. Kataev, arXiv:hep-ph/0602098 (2006); M. Passera, J. Phys. **G31**, 75 (2005).
10. G. Gabrielse *et al.*, Phys. Rev. Lett. **97**, 030802 (2006); Erratum *ibid.* Phys. Rev. Lett. **99**, 039902 (2007); D. Hanneke, S. Fogwell, and G. Gabrielse, Phys. Rev. Lett. **100**, 120801 (2008).
11. R. Bouchendira *et al.*, Phys. Rev. Lett. **106**, 080801 (2011).
12. R. Jackiw and S. Weinberg, Phys. Rev. **D5**, 2396 (1972); G. Altarelli *et al.*, Phys. Lett. **B40**, 415 (1972); I. Bars and M. Yoshimura, Phys. Rev. **D6**, 374 (1972); K. Fujikawa, B.W. Lee, and A.I. Sanda, Phys. Rev. **D6**, 2923 (1972).
13. C. Gnendiger, D. Stöckinger, H. Stöckinger-Kim, Phys. Rev. **D88**, 053005 (2013); A. Czarnecki *et al.*, Phys. Rev. **D67**, 073006 (2003), Erratum *ibid.* Phys. Rev. **D73**, 119901 (2006); S. Heinemeyer, D. Stockinger, and G. Weiglein, Nucl. Phys. **B699**, 103 (2004); T. Gribouk and A. Czarnecki, Phys. Rev. **D72**, 053016 (2005); A. Czarnecki, B. Krause, and W.J. Marciano, Phys. Rev. Lett. **76**, 3267 (1996); A. Czarnecki, B. Krause, and W.J. Marciano, Phys. Rev. **D52**, 2619 (1995); S. Peris, M. Perrottet, and E. de Rafael, Phys. Lett. **B355**, 523 (1995); T. Kukhto *et al.*, Nucl. Phys. **B371**, 567 (1992).
14. G. Degrossi and G.F. Giudice, Phys. Rev. **D58**, 053007 (1998).

## Lepton Particle Listings

 $\mu$ 

15. C. Bouchiat and L. Michel, J. Phys. Radium **22**, 121 (1961);  
M. Gourdin and E. de Rafael, Nucl. Phys. **B10**, 667 (1969).
16. S.J. Brodsky and E. de Rafael, Phys. Rev. **168**, 1620 (1968).
17. M. Davier *et al.*, Eur. Phys. J. **C71**, 1515 (2011).
18. R. Alemany *et al.*, Eur. Phys. J. **C2**, 123 (1998).
19. M. Davier *et al.*, Eur. Phys. J. **C66**, 127 (2010).
20. BABAR Collaboration (B. Aubert *et al.*), Phys. Rev. Lett. **103**, 231801 (2009).
21. K. Hagiwara *et al.*, JPHGB **G38**, 085003 (2011).
22. S. Bodenstern *et al.*, Phys. Rev. **D88**, 014005 (2013).
23. M. Benayoun *et al.*, Eur. Phys. J. **C73**, 2453 (2013).
24. B.Krause, Phys. Lett. **B390**, 392 (1997).
25. J. Bijnens and J. Prades, Mod. Phys. Lett. **A22**, 767 (2007).
26. K. Melnikov and A. Vainshtein, Phys. Rev. **D70**, 113006 (2004).
27. M. Knecht and A. Nyffeler, Phys. Rev. **D65**, 073034 (2002);  
M. Knecht *et al.*, Phys. Rev. Lett. **88**, 071802 (2002).
28. J. Bijnens *et al.*, Nucl. Phys. **B626**, 410 (2002).
29. J. Prades, E. de Rafael, and A. Vainshtein,  
Advanced series on directions in high energy physics 20,  
Editors B.L. Roberts and W. Marciano,  
arXiv:0901.0306 [hep-ph] (2009).
30. J. Hayakawa and T. Kinoshita, Erratum Phys. Rev. **D66**, 019902 (2002).
31. E. de Rafael, Phys. Lett. **B322**, 239 (1994).
32. F. Jegerlehner and R. Szafron, Eur. Phys. J. **C71**, 1632 (2011).
33. P. Fayet, Phys. Rev. **D75**, 115017 (2007).
34. M. Pospelov, Phys. Rev. **D80**, 095002 (2009).
35. D. Tucker-Smith and I. Yavin, Phys. Rev. **D83**, 101702 (R)(2011).

 $\mu$  MAGNETIC MOMENT ANOMALY

The parity-violating decay of muons in a storage ring is observed. The difference frequency  $\omega_a$  between the muon spin precession and the orbital angular frequency ( $e/m_\mu c \langle B \rangle$ ) is measured, as is the free proton NMR frequency  $\omega_p$ , thus determining the ratio  $R = \omega_a/\omega_p$ . Given the magnetic moment ratio  $\lambda = \mu_\mu/\mu_p$  (from hyperfine structure in muonium),  $(g-2)/2 = R/(\lambda - R)$ .

$$\mu_\mu/(e\hbar/2m_\mu) - 1 = (g_\mu - 2)/2$$

VALUE (units $10^{-10}$ )	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>11659208.9 \pm 5.4 \pm 3.3</math></b>	<sup>1</sup> BENNETT	06	MUG2	Average $\mu^+$ and $\mu^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
11659208 $\pm$ 6	BENNETT	04	MUG2	Average $\mu^+$ and $\mu^-$
11659214 $\pm$ 8 $\pm$ 3	BENNETT	04	MUG2	Storage ring
11659203 $\pm$ 6 $\pm$ 5	BENNETT	04	MUG2	Storage ring
11659204 $\pm$ 7 $\pm$ 5	BENNETT	02	MUG2	Storage ring
11659202 $\pm$ 14 $\pm$ 6	BROWN	01	MUG2	Storage ring
11659191 $\pm$ 59	BROWN	00	MUG2	Storage ring
11659100 $\pm$ 110	<sup>2</sup> BAILEY	79	CNTR	Storage ring
11659360 $\pm$ 120	<sup>2</sup> BAILEY	79	CNTR	Storage ring
11659230 $\pm$ 85	<sup>2</sup> BAILEY	79	CNTR	Storage ring
11620000 $\pm$ 5000	CHARPAK	62	CNTR	Storage ring

<sup>1</sup> BENNETT 06 reports  $(g_\mu - 2)/2 = (11659208.0 \pm 5.4 \pm 3.3) \times 10^{-10}$ . We rescaled this value using  $\mu/p$  magnetic moment ratio of 3.183345137(85) from MOHR 08.

<sup>2</sup> BAILEY 79 values recalculated by HUGHES 99 using the COHEN 87  $\mu/p$  magnetic moment. The improved MOHR 99 value does not change the result.

$$(g_{\mu^+} - g_{\mu^-}) / g_{\text{average}}$$

A test of  $CPT$  invariance.

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN
<b><math>-0.11 \pm 0.12</math></b>	BENNETT	04 MUG2
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$-2.6 \pm 1.6$	BAILEY	79 CNTR

 $\mu$  ELECTRIC DIPOLE MOMENT (d)

A nonzero value is forbidden by both  $T$  invariance and  $P$  invariance.

VALUE ( $10^{-19}$ ecm)	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>-0.1 \pm 0.9</math></b>	<sup>1</sup> BENNETT	09	MUG2	$\pm$ Storage ring
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.1 \pm 1.0$	BENNETT	09	MUG2	$+$ Storage ring
$-0.1 \pm 0.7$	BENNETT	09	MUG2	$-$ Storage ring
$-3.7 \pm 3.4$	<sup>2</sup> BAILEY	78	CNTR	$\pm$ Storage ring
$8.6 \pm 4.5$	BAILEY	78	CNTR	$+$ Storage ring
$0.8 \pm 4.3$	BAILEY	78	CNTR	$-$ Storage ring

<sup>1</sup> This is the combination of the two BENNETT 09 results quoted here separately for  $\mu^+$  and  $\mu^-$ . BENNETT 09 uses the convention  $d = 1/2 \cdot (d_{\mu^-} - d_{\mu^+})$ .

<sup>2</sup> This is the combination of the two BAILEY 78 results quoted here separately for  $\mu^+$  and  $\mu^-$ . BAILEY 78 uses the convention  $d = 1/2 \cdot (d_{\mu^+} - d_{\mu^-})$  and reports  $3.7 \pm 3.4$ . We convert their result to use the same convention as BENNETT 09.

MUON-ELECTRON CHARGE RATIO ANOMALY  $q_{\mu^+}/q_{e^-} + 1$ 

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>(1.1 \pm 2.1) \times 10^{-9}</math></b>	<sup>1</sup> MEYER	00	CNTR	$+$ 1s–2s muonium interval

<sup>1</sup> MEYER 00 measure the 1s–2s muonium interval, and then interpret the result in terms of muon-electron charge ratio  $q_{\mu^+}/q_{e^-}$ .

 $\mu^-$  DECAY MODES

$\mu^+$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $e^- \bar{\nu}_e \nu_\mu$	$\approx 100\%$	
$\Gamma_2$ $e^- \bar{\nu}_e \nu_\mu \gamma$	[a] $(1.4 \pm 0.4)\%$	
$\Gamma_3$ $e^- \bar{\nu}_e \nu_\mu e^+ e^-$	[b] $(3.4 \pm 0.4) \times 10^{-5}$	

## Lepton Family number (LF) violating modes

Mode	LF	$\Gamma_i/\Gamma$	Confidence level
$\Gamma_4$ $e^- \nu_e \bar{\nu}_\mu$	LF	[c] $< 1.2\%$	90%
$\Gamma_5$ $e^- \gamma$	LF	$< 5.7 \times 10^{-13}$	90%
$\Gamma_6$ $e^- e^+ e^-$	LF	$< 1.0 \times 10^{-12}$	90%
$\Gamma_7$ $e^- 2\gamma$	LF	$< 7.2 \times 10^{-11}$	90%

[a] This only includes events with the  $\gamma$  energy  $> 10$  MeV. Since the  $e^- \bar{\nu}_e \nu_\mu$  and  $e^- \bar{\nu}_e \nu_\mu \gamma$  modes cannot be clearly separated, we regard the latter mode as a subset of the former.

[b] See the Particle Listings below for the energy limits used in this measurement.

[c] A test of additive vs. multiplicative lepton family number conservation.

 $\mu^-$  BRANCHING RATIOS

$\Gamma(e^- \bar{\nu}_e \nu_\mu \gamma)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
<b><math>0.014 \pm 0.004</math></b>			CRITTENDEN	61	CNTR $\gamma$ KE $> 10$ MeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
	862		BOGART	67	CNTR $\gamma$ KE $> 14.5$ MeV	
$0.0033 \pm 0.0013$			CRITTENDEN	61	CNTR $\gamma$ KE $> 20$ MeV	
	27		ASHKIN	59	CNTR	

$\Gamma(e^- \bar{\nu}_e \nu_\mu e^+ e^-)/\Gamma_{\text{total}}$	VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_3/\Gamma$
<b><math>3.4 \pm 0.2 \pm 0.3</math></b>		7443	<sup>1</sup> BERTL	85	SPEC	$+$ SINDRUM	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
$2.2 \pm 1.5$		7	<sup>2</sup> CRITTENDEN	61	HLBC	$+$ $E(e^+ e^-) > 10$ MeV	
$2$		1	<sup>3</sup> GUREVICH	60	EMUL	$+$	
$1.5 \pm 1.0$		3	<sup>4</sup> LEE	59	HBC	$+$	

<sup>1</sup> BERTL 85 has transverse momentum cut  $p_T > 17$  MeV/c. Systematic error was increased by us.

<sup>2</sup> CRITTENDEN 61 count only those decays where total energy of either ( $e^+$ ,  $e^-$ ) combination is  $> 10$  MeV.

<sup>3</sup> GUREVICH 60 interpret their event as either virtual or real photon conversion.  $e^+$  and  $e^-$  energies not measured.

<sup>4</sup> In the three LEE 59 events, the sum of energies  $E(e^+) + E(e^-) + E(e^+)$  was 51 MeV, 55 MeV, and 33 MeV.

$\Gamma(e^- \nu_e \bar{\nu}_\mu) / \Gamma_{\text{total}}$   $\Gamma_4 / \Gamma$   
Forbidden by the additive conservation law for lepton family number. A multiplicative law predicts this branching ratio to be 1/2. For a review see NEMETHY 81.

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< 0.012	90	<sup>1</sup> FREEDMAN	93	CNTR	+ $\nu$ oscillation search
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< 0.018	90	KRAKAUER	918	CALO	+
< 0.05	90	<sup>2</sup> BERGSMA	83	CALO	$\bar{\nu}_\mu e \rightarrow \mu^- \bar{\nu}_e$
< 0.09	90	JONKER	80	CALO	See BERGSMA 83
-0.001 ± 0.061		WILLIS	80	CNTR	+
0.13 ± 0.15		BLIETSCHAU	78	HLBC	± Avg. of 4 values
< 0.25	90	EICHEN	73	HLBC	+

<sup>1</sup> FREEDMAN 93 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton family number violation.

<sup>2</sup> BERGSMA 83 gives a limit on the inverse muon decay cross-section ratio  $\sigma(\bar{\nu}_\mu e^- \rightarrow \mu^- \bar{\nu}_e) / \sigma(\nu_\mu e^- \rightarrow \mu^- \nu_e)$ , which is essentially equivalent to  $\Gamma(e^- \nu_e \bar{\nu}_\mu) / \Gamma_{\text{total}}$  for small values like that quoted.

$\Gamma(e^- \gamma) / \Gamma_{\text{total}}$   $\Gamma_5 / \Gamma$   
Forbidden by lepton family number conservation.

VALUE (units $10^{-11}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< 0.057	90	ADAM	138	SPEC	+ MEG at PSI
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< 0.24	90	ADAM	11	SPEC	+ MEG at PSI
< 2.8	90	ADAM	10	SPEC	+ MEG at PSI
< 1.2	90	AHMED	02	SPEC	+ MEGA
< 1.2	90	BROOKS	99	SPEC	+ LAMPF
< 4.9	90	BOLTON	88	CBOX	+ LAMPF
< 100	90	AZUELOS	83	CNTR	+ TRIUMF
< 17	90	KINNISON	82	SPEC	+ LAMPF
< 100	90	SCHAAF	80	ELEC	+ SIN

$\Gamma(e^- e^+ e^-) / \Gamma_{\text{total}}$   $\Gamma_6 / \Gamma$   
Forbidden by lepton family number conservation.

VALUE (units $10^{-12}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< 1.0	90	<sup>1</sup> BELLGARDT	88	SPEC	+ SINDRUM
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< 36	90	BARANOV	91	SPEC	+ ARES
< 35	90	BOLTON	88	CBOX	+ LAMPF
< 2.4	90	<sup>1</sup> BERTL	85	SPEC	+ SINDRUM
< 160	90	<sup>1</sup> BERTL	84	SPEC	+ SINDRUM
< 130	90	<sup>1</sup> BOLTON	84	CNTR	+ LAMPF

<sup>1</sup> These experiments assume a constant matrix element.

$\Gamma(e^- 2\gamma) / \Gamma_{\text{total}}$   $\Gamma_7 / \Gamma$   
Forbidden by lepton family number conservation.

VALUE (units $10^{-11}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< 7.2	90	BOLTON	88	CBOX	+ LAMPF
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< 840	90	<sup>1</sup> AZUELOS	83	CNTR	+ TRIUMF
< 5000	90	<sup>2</sup> BOWMAN	78	CNTR	DEPOMMIER 77 data

<sup>1</sup> AZUELOS 83 uses the phase space distribution of BOWMAN 78.

<sup>2</sup> BOWMAN 78 assumes an interaction Lagrangian local on the scale of the inverse  $\mu$  mass.

### LIMIT ON $\mu^- \rightarrow e^-$ CONVERSION

Forbidden by lepton family number conservation.

$\sigma(\mu^- {}^{32}\text{S} \rightarrow e^- {}^{32}\text{S}) / \sigma(\mu^- {}^{32}\text{S} \rightarrow \nu_\mu {}^{32}\text{P}^*)$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< $7 \times 10^{-11}$	90	BADERT...	80	STRC	SIN
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< $4 \times 10^{-10}$	90	BADERT...	77	STRC	SIN

$\sigma(\mu^- \text{Cu} \rightarrow e^- \text{Cu}) / \sigma(\mu^- \text{Cu} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< $1.6 \times 10^{-8}$	90	BRYMAN	72	SPEC	

$\sigma(\mu^- \text{Ti} \rightarrow e^- \text{Ti}) / \sigma(\mu^- \text{Ti} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< $4.3 \times 10^{-12}$	90	<sup>1</sup> DOHMEN	93	SPEC	SINDRUM II
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< $4.6 \times 10^{-12}$	90	AHMAD	88	TPC	TRIUMF
< $1.6 \times 10^{-11}$	90	BRYMAN	85	TPC	TRIUMF

<sup>1</sup> DOHMEN 93 assumes  $\mu^- \rightarrow e^-$  conversion leaves the nucleus in its ground state, a process enhanced by coherence and expected to dominate.

$\sigma(\mu^- \text{Pb} \rightarrow e^- \text{Pb}) / \sigma(\mu^- \text{Pb} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< $4.6 \times 10^{-11}$	90	HONECKER	96	SPEC	SINDRUM II
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< $4.9 \times 10^{-10}$	90	AHMAD	88	TPC	TRIUMF

$\sigma(\mu^- \text{Au} \rightarrow e^- \text{Au}) / \sigma(\mu^- \text{Au} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< $7 \times 10^{-13}$	90	BERTL	06	SPEC	- SINDRUM II

### LIMIT ON $\mu^- \rightarrow e^+$ CONVERSION

Forbidden by total lepton number conservation.

$\sigma(\mu^- {}^{32}\text{S} \rightarrow e^+ {}^{32}\text{Si}^*) / \sigma(\mu^- {}^{32}\text{S} \rightarrow \nu_\mu {}^{32}\text{P}^*)$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< $9 \times 10^{-10}$	90	BADERT...	80	STRC	SIN
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< $1.5 \times 10^{-9}$	90	BADERT...	78	STRC	SIN

$\sigma(\mu^- {}^{127}\text{I} \rightarrow e^+ {}^{127}\text{Sb}^*) / \sigma(\mu^- {}^{127}\text{I} \rightarrow \text{anything})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< $3 \times 10^{-10}$	90	<sup>1</sup> ABELA	80	CNTR	Radiochemical tech.

<sup>1</sup> ABELA 80 is upper limit for  $\mu^- e^+$  conversion leading to particle-stable states of  ${}^{127}\text{Sb}$ . Limit for total conversion rate is higher by a factor less than 4 (G. Backenstoss, private communication).

$\sigma(\mu^- \text{Cu} \rightarrow e^+ \text{Co}) / \sigma(\mu^- \text{Cu} \rightarrow \nu_\mu \text{Ni})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
< $2.6 \times 10^{-8}$	90	BRYMAN	72	SPEC	
< $2.2 \times 10^{-7}$	90	CONFORTO	62	OSPK	

$\sigma(\mu^- \text{Ti} \rightarrow e^+ \text{Ca}) / \sigma(\mu^- \text{Ti} \rightarrow \text{capture})$

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<3.6 × 10 <sup>-11</sup>	90	1	<sup>1,2</sup> KAULARD	98	SPEC	— SINDRUM II
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
<1.7 × 10 <sup>-12</sup>	90	1	<sup>2,3</sup> KAULARD	98	SPEC	— SINDRUM II
<4.3 × 10 <sup>-12</sup>	90		<sup>3</sup> DOHMEN	93	SPEC	SINDRUM II
<8.9 × 10 <sup>-11</sup>	90		<sup>1</sup> DOHMEN	93	SPEC	SINDRUM II
<1.7 × 10 <sup>-10</sup>	90		<sup>4</sup> AHMAD	88	TPC	TRIUMF

<sup>1</sup> This limit assumes a giant resonance excitation of the daughter Ca nucleus (mean energy and width both 20 MeV).

<sup>2</sup> KAULARD 98 obtained these same limits using the unified classical analysis of FELD-MAN 98.

<sup>3</sup> This limit assumes the daughter Ca nucleus is left in the ground state. However, the probability of this is unknown.

<sup>4</sup> Assuming a giant-resonance-excitation model.

### LIMIT ON MUONIUM $\rightarrow$ ANTIMUONIUM CONVERSION

Forbidden by lepton family number conservation.

$R_g = G_C / G_F$

The effective Lagrangian for the  $\mu^+ e^- \rightarrow \mu^- e^+$  conversion is assumed to be

$$\mathcal{L} = 2^{-1/2} G_C [\bar{\psi}_\mu \gamma_\lambda (1 - \gamma_5) \psi_e] [\bar{\psi}_\mu \gamma_\lambda (1 - \gamma_5) \psi_e] + \text{h.c.}$$

The experimental result is then an upper limit on  $G_C / G_F$ , where  $G_F$  is the Fermi coupling constant.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
< 0.0030	90	1	<sup>1</sup> WILLMANN	99	SPEC	+ $\mu^+$ at 26 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •						
< 0.14	90	1	<sup>2</sup> GORDEEV	97	SPEC	+ JINR phasotron
< 0.018	90	0	<sup>3</sup> ABELA	96	SPEC	+ $\mu^+$ at 24 MeV
< 6.9	90		NI	93	CBOX	LAMPF
< 0.16	90		MATTHIAS	91	SPEC	LAMPF
< 0.29	90		HUBER	90B	CNTR	TRIUMF
< 20	95		BEER	86	CNTR	TRIUMF
< 42	95		MARSHALL	82	CNTR	

<sup>1</sup> WILLMANN 99 quote both probability  $P_{M\bar{M}} < 8.3 \times 10^{-11}$  at 90%CL in a 0.1 T field and  $R_g = G_C / G_F$ .

<sup>2</sup> GORDEEV 97 quote limits on both  $f = G_{MM} / G_F$  and the probability  $W_{MM} < 4.7 \times 10^{-7}$  (90% CL).

<sup>3</sup> ABELA 96 quote both probability  $P_{M\bar{M}} < 8 \times 10^{-9}$  at 90% CL and  $R_g = G_C / G_F$ .

### MUON DECAY PARAMETERS

Revised September 2013 by W. Fetscher and H.-J. Gerber (ETH Zürich).

**Introduction:** All measurements in direct muon decay,  $\mu^- \rightarrow e^- + 2 \text{ neutrals}$ , and its inverse,  $\nu_\mu + e^- \rightarrow \mu^- + \text{neutral}$ , are successfully described by the "V-A interaction," which is a particular case of a local, derivative-free, lepton-number-conserving,

# Lepton Particle Listings

$\mu$

four-fermion interaction [1]. As shown below, within this framework, the Standard Model assumptions, such as the  $V-A$  form and the nature of the neutrals ( $\nu_\mu$  and  $\bar{\nu}_e$ ), and hence the doublet assignments ( $\nu_e e^-$ ) $_L$  and ( $\nu_\mu \mu^-$ ) $_L$ , have been determined from experiments [2,3]. All considerations on muon decay are valid for the leptonic tau decays  $\tau \rightarrow \ell + \nu_\tau + \bar{\nu}_e$  with the replacements  $m_\mu \rightarrow m_\tau$ ,  $m_e \rightarrow m_\ell$ .

**Parameters:** The differential decay probability to obtain an  $e^\pm$  with (reduced) energy between  $x$  and  $x + dx$ , emitted in the direction  $\hat{\mathbf{x}}_3$  at an angle between  $\vartheta$  and  $\vartheta + d\vartheta$  with respect to the muon polarization vector  $\mathbf{P}_\mu$ , and with its spin parallel to the arbitrary direction  $\hat{\boldsymbol{\zeta}}$ , neglecting radiative corrections, is given by

$$\begin{aligned} \frac{d^2\Gamma}{dx d\cos\vartheta} &= \frac{m_\mu}{4\pi^3} W_{e\mu}^4 G_F^2 \sqrt{x^2 - x_0^2} \\ &\times (F_{IS}(x) \pm P_\mu \cos\vartheta F_{AS}(x)) \\ &\times \left[ 1 + \hat{\boldsymbol{\zeta}} \cdot \mathbf{P}_e(x, \vartheta) \right] . \end{aligned} \quad (1)$$

Here,  $W_{e\mu} = \max(E_e) = (m_\mu^2 + m_e^2)/2m_\mu$  is the maximum  $e^\pm$  energy,  $x = E_e/W_{e\mu}$  is the reduced energy,  $x_0 = m_e/W_{e\mu} = 9.67 \times 10^{-3}$ , and  $P_\mu = |\mathbf{P}_\mu|$  is the degree of muon polarization.  $\hat{\boldsymbol{\zeta}}$  is the direction in which a perfect polarization-sensitive electron detector is most sensitive. The isotropic part of the spectrum,  $F_{IS}(x)$ , the anisotropic part  $F_{AS}(x)$ , and the electron polarization,  $\mathbf{P}_e(x, \vartheta)$ , may be parametrized by the Michel parameter  $\rho$  [1], by  $\eta$  [4], by  $\xi$  and  $\delta$  [5,6], *etc.* These are bilinear combinations of the coupling constants  $g_{e\mu}^2$ , which occur in the matrix element (given below).

If the masses of the neutrinos as well as  $x_0^2$  are neglected, the energy and angular distribution of the electron in the rest frame of a muon ( $\mu^\pm$ ) measured by a polarization insensitive detector, is given by

$$\begin{aligned} \frac{d^2\Gamma}{dx d\cos\vartheta} &\sim x^2 \cdot \left\{ 3(1-x) + \frac{2\rho}{3}(4x-3) + 3\eta x_0(1-x)/x \right. \\ &\quad \left. \pm P_\mu \cdot \xi \cdot \cos\vartheta \left[ 1 - x + \frac{2\delta}{3}(4x-3) \right] \right\} . \end{aligned} \quad (2)$$

Here,  $\vartheta$  is the angle between the electron momentum and the muon spin, and  $x \equiv 2E_e/m_\mu$ . For the Standard Model coupling, we obtain  $\rho = \xi\delta = 3/4$ ,  $\xi = 1$ ,  $\eta = 0$  and the differential decay rate is

$$\frac{d^2\Gamma}{dx d\cos\vartheta} = \frac{G_F^2 m_\mu^5}{192\pi^3} [3 - 2x \pm P_\mu \cos\vartheta(2x-1)] x^2 . \quad (3)$$

The coefficient in front of the square bracket is the total decay rate.

If only the neutrino masses are neglected, and if the  $e^\pm$  polarization is detected, then the functions in Eq. (1) become

$$\begin{aligned} F_{IS}(x) &= x(1-x) + \frac{2}{9} \rho(4x^2 - 3x - x_0^2) + \eta \cdot x_0(1-x) \\ F_{AS}(x) &= \frac{1}{3} \xi \sqrt{x^2 - x_0^2} \\ &\quad \times [1 - x + \frac{2}{3} \delta(4x - 3 + (\sqrt{1 - x_0^2} - 1))] \\ \mathbf{P}_e(x, \vartheta) &= P_{T1} \cdot \hat{\mathbf{x}}_1 + P_{T2} \cdot \hat{\mathbf{x}}_2 + P_L \cdot \hat{\mathbf{x}}_3 . \end{aligned} \quad (4)$$

Here  $\hat{\mathbf{x}}_1$ ,  $\hat{\mathbf{x}}_2$ , and  $\hat{\mathbf{x}}_3$  are orthogonal unit vectors defined as follows:

$$\begin{aligned} \hat{\mathbf{x}}_3 &\text{ is along the } e \text{ momentum } \mathbf{p}_e \\ \frac{\hat{\mathbf{x}}_3 \times \mathbf{P}_\mu}{|\hat{\mathbf{x}}_3 \times \mathbf{P}_\mu|} &= \hat{\mathbf{x}}_2 \text{ is transverse to } \mathbf{p}_e \text{ and perpendicular} \\ &\text{to the “decay plane”} \\ \hat{\mathbf{x}}_2 \times \hat{\mathbf{x}}_3 &= \hat{\mathbf{x}}_1 \text{ is transverse to the } \mathbf{p}_e \text{ and in the} \\ &\text{“decay plane.”} \end{aligned}$$

The components of  $\mathbf{P}_e$  then are given by

$$\begin{aligned} P_{T1}(x, \vartheta) &= P_\mu \sin\vartheta \cdot F_{T1}(x) / (F_{IS}(x) \pm P_\mu \cos\vartheta \cdot F_{AS}(x)) \\ P_{T2}(x, \vartheta) &= P_\mu \sin\vartheta \cdot F_{T2}(x) / (F_{IS}(x) \pm P_\mu \cos\vartheta \cdot F_{AS}(x)) \\ P_L(x, \vartheta) &= \left( \pm F_{IP}(x) + P_\mu \cos\vartheta \right. \\ &\quad \left. \times F_{AP}(x) \right) / (F_{IS}(x) \pm P_\mu \cos\vartheta \cdot F_{AS}(x)) , \end{aligned}$$

where

$$\begin{aligned} F_{T1}(x) &= \frac{1}{12} \left\{ -2 \left[ \xi'' + 12(\rho - \frac{3}{4}) \right] (1-x)x_0 \right. \\ &\quad \left. - 3\eta(x^2 - x_0^2) + \eta''(-3x^2 + 4x - x_0^2) \right\} \\ F_{T2}(x) &= \frac{1}{3} \sqrt{x^2 - x_0^2} \left\{ 3 \frac{\alpha'}{A} (1-x) + 2 \frac{\beta'}{A} \sqrt{1 - x_0^2} \right\} \\ F_{IP}(x) &= \frac{1}{54} \sqrt{x^2 - x_0^2} \left\{ 9\xi' \left( -2x + 2 + \sqrt{1 - x_0^2} \right) \right. \\ &\quad \left. + 4\xi(\delta - \frac{3}{4})(4x - 4 + \sqrt{1 - x_0^2}) \right\} \\ F_{AP}(x) &= \frac{1}{6} \left\{ \xi''(2x^2 - x - x_0^2) + 4(\rho - \frac{3}{4})(4x^2 - 3x - x_0^2) \right. \\ &\quad \left. + 2\eta''(1-x)x_0 \right\} . \end{aligned} \quad (5)$$

For the experimental values of the parameters  $\rho$ ,  $\xi$ ,  $\xi'$ ,  $\xi''$ ,  $\delta$ ,  $\eta$ ,  $\eta''$ ,  $\alpha/A$ ,  $\beta/A$ ,  $\alpha'/A$ ,  $\beta'/A$ , which are not all independent, see the Data Listings below. Experiments in the past have also been analyzed using the parameters  $a$ ,  $b$ ,  $c$ ,  $a'$ ,  $b'$ ,  $c'$ ,  $\alpha/A$ ,  $\beta/A$ ,  $\alpha'/A$ ,  $\beta'/A$  (and  $\eta = (\alpha - 2\beta)/2A$ ), as defined by Kinoshita and Sirlin [5,6]. They serve as a model-independent summary of all possible measurements on the decay electron (see Listings below). The relations between the two sets of parameters are

$$\begin{aligned} \rho - \frac{3}{4} &= \frac{3}{4}(-a + 2c)/A , \\ \eta &= (\alpha - 2\beta)/A , \\ \eta'' &= (3\alpha + 2\beta)/A , \\ \delta - \frac{3}{4} &= \frac{9}{4} \cdot \frac{(a' - 2c')/A}{1 - [a + 3a' + 4(b + b') + 6c - 14c']/A} , \\ 1 - \xi \frac{\delta}{\rho} &= 4 \frac{[(b + b') + 2(c - c')]/A}{1 - (a - 2c)/A} , \\ 1 - \xi' &= [(a + a') + 4(b + b') + 6(c + c')]/A , \\ 1 - \xi'' &= (-2a + 20c)/A , \end{aligned}$$

where

$$A = a + 4b + 6c . \quad (6)$$

The differential decay probability to obtain a *left-handed*  $\nu_e$  with (reduced) energy between  $y$  and  $y + dy$ , neglecting radiative



corrections as well as the masses of the electron and of the neutrinos, is given by [7]

$$\frac{d\Gamma}{dy} = \frac{m_\mu^5 G_F^2}{16\pi^3} \cdot Q_L^{\nu_e} \cdot y^2 \left\{ (1-y) - \omega_L \cdot (y - \frac{3}{4}) \right\}. \quad (7)$$

Here,  $y = 2 E_{\nu_e}/m_\mu$ ,  $Q_L^{\nu_e}$  and  $\omega_L$  are parameters.  $\omega_L$  is the neutrino analog of the spectral shape parameter  $\rho$  of Michel. Since in the Standard Model,  $Q_L^{\nu_e} = 1$ ,  $\omega_L = 0$ , the measurement of  $d\Gamma/dy$  has allowed a null-test of the Standard Model (see Listings below).

**Matrix element:** All results in direct muon decay (energy spectra of the electron and of the neutrinos, polarizations, and angular distributions), and in inverse muon decay (the reaction cross section) at energies well below  $m_W c^2$ , may be parametrized in terms of amplitudes  $g_{\varepsilon\mu}^\gamma$  and the Fermi coupling constant  $G_F$ , using the matrix element

$$\frac{4G_F}{\sqrt{2}} \sum_{\substack{\gamma=S,V,T \\ \varepsilon,\mu=R,L}} g_{\varepsilon\mu}^\gamma \langle \bar{e}_\varepsilon | \Gamma^\gamma | (\nu_e)_n \rangle \langle (\bar{\nu}_\mu)_m | \Gamma_\gamma | \mu_\mu \rangle. \quad (8)$$

We use the notation of Fetscher *et al.* [2], who in turn use the sign conventions and definitions of Scheck [8]. Here,  $\gamma = S, V, T$  indicates a scalar, vector, or tensor interaction; and  $\varepsilon, \mu = R, L$  indicate a right- or left-handed chirality of the electron or muon. The chiralities  $n$  and  $m$  of the  $\nu_e$  and  $\bar{\nu}_\mu$  are then determined by the values of  $\gamma, \varepsilon$ , and  $\mu$ . The particles are represented by fields of definite chirality [9].

As shown by Langacker and London [10], explicit lepton-number nonconservation still leads to a matrix element equivalent to Eq. (8). They conclude that it is not possible, even in principle, to test lepton-number conservation in (leptonic) muon decay if the final neutrinos are massless and are not observed.

The ten complex amplitudes  $g_{\varepsilon\mu}^\gamma$  ( $g_{RR}^T$  and  $g_{LL}^T$  are identically zero) and  $G_F$  constitute 19 independent (real) parameters to be determined by experiment. The Standard Model interaction corresponds to one single amplitude  $g_{LL}^V$  being unity and all the others being zero.

The (direct) muon decay experiments are compatible with an arbitrary mix of the scalar and vector amplitudes  $g_{LL}^S$  and  $g_{LL}^V$  – in the extreme even with purely scalar  $g_{LL}^S = 2$ ,  $g_{LL}^V = 0$ . The decision in favour of the Standard Model comes from the quantitative observation of inverse muon decay, which would be forbidden for pure  $g_{LL}^S$  [2].

**Experimental determination of V-A:** In order to determine the amplitudes  $g_{\varepsilon\mu}^\gamma$  uniquely from experiment, the following set of equations, where the left-hand sides represent experimental results, has to be solved.

$$\begin{aligned} a &= 16(|g_{RL}^V|^2 + |g_{LR}^V|^2) + |g_{RL}^S + 6g_{RL}^T|^2 + |g_{LR}^S + 6g_{LR}^T|^2 \\ a' &= 16(|g_{RL}^V|^2 - |g_{LR}^V|^2) + |g_{RL}^S + 6g_{RL}^T|^2 - |g_{LR}^S + 6g_{LR}^T|^2 \\ \alpha &= 8\text{Re} \left\{ g_{RL}^V (g_{RL}^{S*} + 6g_{RL}^{T*}) + g_{LR}^V (g_{LR}^{S*} + 6g_{LR}^{T*}) \right\} \\ \alpha' &= 8\text{Im} \left\{ g_{LR}^V (g_{RL}^{S*} + 6g_{RL}^{T*}) - g_{RL}^V (g_{LR}^{S*} + 6g_{LR}^{T*}) \right\} \\ b &= 4(|g_{RR}^V|^2 + |g_{LL}^V|^2) + |g_{RR}^S|^2 + |g_{LL}^S|^2 \end{aligned}$$

$$b' = 4(|g_{RR}^V|^2 - |g_{LL}^V|^2) + |g_{RR}^S|^2 - |g_{LL}^S|^2$$

$$\beta = -4\text{Re} \left\{ g_{RR}^V g_{LL}^{S*} + g_{LL}^V g_{RR}^{S*} \right\}$$

$$\beta' = 4\text{Im} \left\{ g_{RR}^V g_{LL}^{S*} - g_{LL}^V g_{RR}^{S*} \right\}$$

$$c = \frac{1}{2} \left\{ |g_{RL}^S - 2g_{RL}^T|^2 + |g_{LR}^S - 2g_{LR}^T|^2 \right\}$$

$$c' = \frac{1}{2} \left\{ |g_{RL}^S - 2g_{RL}^T|^2 - |g_{LR}^S - 2g_{LR}^T|^2 \right\}$$

and

$$Q_L^{\nu_e} = 1 - \left\{ \frac{1}{4}|g_{LR}^S|^2 + \frac{1}{4}|g_{LL}^S|^2 + |g_{RR}^V|^2 + |g_{RL}^V|^2 + 3|g_{LR}^T|^2 \right\}$$

$$\omega_L = \frac{3}{4} \frac{\{|g_{RR}^S|^2 + 4|g_{LR}^V|^2 + |g_{RL}^S + 2g_{RL}^T|^2\}}{|g_{RL}^S|^2 + |g_{RR}^S|^2 + 4|g_{LL}^V|^2 + 4|g_{LR}^V|^2 + 12|g_{RL}^T|^2}.$$

It has been noted earlier by C. Jarlskog [11], that certain experiments observing the decay electron are especially informative if they yield the  $V$ - $A$  values. The complete solution is now found as follows. Fetscher *et al.* [2] introduced four probabilities  $Q_{\varepsilon\mu}(\varepsilon, \mu = R, L)$  for the decay of a  $\mu$ -handed muon into an  $\varepsilon$ -handed electron, and showed that there exist upper bounds on  $Q_{RR}$ ,  $Q_{LR}$ , and  $Q_{RL}$ , and a lower bound on  $Q_{LL}$ . These probabilities are given in terms of the  $g_{\varepsilon\mu}^\gamma$ 's by

$$Q_{\varepsilon\mu} = \frac{1}{4}|g_{\varepsilon\mu}^S|^2 + |g_{\varepsilon\mu}^V|^2 + 3(1 - \delta_{\varepsilon\mu})|g_{\varepsilon\mu}^T|^2, \quad (9)$$

where  $\delta_{\varepsilon\mu} = 1$  for  $\varepsilon = \mu$ , and  $\delta_{\varepsilon\mu} = 0$  for  $\varepsilon \neq \mu$ . They are related to the parameters  $a$ ,  $b$ ,  $c$ ,  $a'$ ,  $b'$ , and  $c'$  by

$$\begin{aligned} Q_{RR} &= 2(b + b')/A, \\ Q_{LR} &= [(a - a') + 6(c - c')]/2A, \\ Q_{RL} &= [(a + a') + 6(c + c')]/2A, \\ Q_{LL} &= 2(b - b')/A, \end{aligned} \quad (10)$$

with  $A = 16$ . In the Standard Model,  $Q_{LL} = 1$  and the others are zero.

Since the upper bounds on  $Q_{RR}$ ,  $Q_{LR}$ , and  $Q_{RL}$  are found to be small, and since the helicity of the  $\nu_\mu$  in pion decay is known from experiment [12,13] to very high precision to be  $-1$  [14], the cross section  $S$  of *inverse* muon decay, normalized to the  $V$ - $A$  value, yields [2]

$$|g_{LL}^S|^2 \leq 4(1 - S) \quad (11)$$

and

$$|g_{LL}^V|^2 = S. \quad (12)$$

Thus the Standard Model assumption of a pure  $V$ - $A$  leptonic charged weak interaction of  $e$  and  $\mu$  is derived (within errors) from experiments at energies far below mass of the  $W^\pm$ : Eq. (12) gives a lower limit for  $V$ - $A$ , and Eqs. (9) and (11) give upper limits for the other four-fermion interactions. The existence of such upper limits may also be seen from  $Q_{RR} + Q_{RL} = (1 - \xi')/2$  and  $Q_{RR} + Q_{LR} = \frac{1}{2}(1 + \xi/3 - 16 \xi \delta/9)$ . Table 1 gives the current experimental limits on the magnitudes of the  $g_{\varepsilon\mu}^\gamma$ 's. More stringent limits on the six coupling constants  $g_{LR}^S$ ,  $g_{LR}^V$ ,

# Lepton Particle Listings

$\mu$

$g_{LR}^T$ ,  $g_{RL}^S$ ,  $g_{RL}^V$ , and  $g_{RL}^T$  have been derived from upper limits on the neutrino mass [18]. Limits on the “charge retention” coordinates, as used in the older literature (e.g., Ref. 19), are given by Burkard *et al.* [20].

**Table 1.** Coupling constants  $g_{\varepsilon\mu}^\gamma$  and some combinations of them. Ninety-percent confidence level experimental limits. The limits on  $|g_{LL}^S|$  and  $|g_{LL}^V|$  are from Ref. 15, and the others from a general analysis of muon decay measurements. Top three rows: Ref. 22, fourth row: Ref. 16, next three rows: Ref. 17, last row: Ref. 21. The experimental uncertainty on the muon polarization in pion decay is included. Note that, by definition,  $|g_{\varepsilon\mu}^S| \leq 2$ ,  $|g_{\varepsilon\mu}^V| \leq 1$  and  $|g_{\varepsilon\mu}^T| \leq 1/\sqrt{3}$ .

$ g_{RR}^S  < 0.035$	$ g_{RR}^V  < 0.017$	$ g_{RR}^T  \equiv 0$
$ g_{LR}^S  < 0.050$	$ g_{LR}^V  < 0.023$	$ g_{LR}^T  < 0.015$
$ g_{RL}^S  < 0.420$	$ g_{RL}^V  < 0.105$	$ g_{RL}^T  < 0.105$
$ g_{LL}^S  < 0.550$	$ g_{LL}^V  > 0.960$	$ g_{LL}^T  \equiv 0$
$ g_{LR}^S + 6g_{LR}^T  < 0.143$	$ g_{RL}^S + 6g_{RL}^T  < 0.418$	
$ g_{LR}^S + 2g_{LR}^T  < 0.108$	$ g_{RL}^S + 2g_{RL}^T  < 0.417$	
$ g_{LR}^S - 2g_{LR}^T  < 0.070$	$ g_{RL}^S - 2g_{RL}^T  < 0.418$	
$Q_{RR} + Q_{LR} < 8.2 \times 10^{-4}$		

## References

1. L. Michel, Proc. Phys. Soc. **A63**, 514 (1950).
2. W. Fetscher, H.-J. Gerber, and K.F. Johnson, Phys. Lett. **B173**, 102 (1986).
3. P. Langacker, Comm. Nucl. Part. Phys. **19**, 1 (1989).
4. C. Bouchiat and L. Michel, Phys. Rev. **106**, 170 (1957).
5. T. Kinoshita and A. Sirlin, Phys. Rev. **107**, 593 (1957).
6. T. Kinoshita and A. Sirlin, Phys. Rev. **108**, 844 (1957).
7. W. Fetscher, Phys. Rev. **D49**, 5945 (1994).
8. F. Scheck, in *Electroweak and Strong Interactions* (Springer Verlag, 1996).
9. K. Mursula and F. Scheck, Nucl. Phys. **B253**, 189 (1985).
10. P. Langacker and D. London, Phys. Rev. **D39**, 266 (1989).
11. C. Jarlskog, Nucl. Phys. **75**, 659 (1966).
12. A. Jodidio *et al.*, Phys. Rev. **D34**, 1967 (1986);  
A. Jodidio *et al.*, Phys. Rev. **D37**, 237 (1988).
13. L.Ph. Roesch *et al.*, Helv. Phys. Acta **55**, 74 (1982).
14. W. Fetscher, Phys. Lett. **140B**, 117 (1984).
15. S.R. Mishra *et al.*, Phys. Lett. **B252**, 170 (1990);  
S.R. Mishra, private communication;  
See also P. Vilain *et al.*, Phys. Lett. **B364**, 121 (1995).
16. R.P. MacDonald *et al.*, Phys. Rev. **D78**, 032010 (2008).
17. C.A. Gagliardi, R.E. Tribble, and N.J. Williams, Phys. Rev. **D72**, 073002 (2005).
18. G. Prézeau and A. Kurylov, Phys. Rev. Lett. **95**, 101802 (2005).
19. S.E. Derenzo, Phys. Rev. **181**, 1854 (1969).
20. H. Burkard *et al.*, Phys. Lett. **160B**, 343 (1985).
21. R. Bayes *et al.*, Phys. Rev. Lett. **106**, 041804 (2011).
22. A. Hillairet *et al.*, Phys. Rev. **D85**, 092013 (2012).

## $\mu$ DECAY PARAMETERS

### $\rho$ PARAMETER

(V–A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.74979 ± 0.00026 OUR AVERAGE</b>					
0.74977 ± 0.00012 ± 0.00023		<sup>1</sup> BAYES	11	TWST	+ Surface $\mu^+$
0.7518 ± 0.0026		DERENZO	69	RVUE	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.75014 ± 0.00017 ± 0.00045		<sup>2</sup> MACDONALD	08	TWST	+ Surface $\mu^+$
0.75080 ± 0.00032 ± 0.00100	6G	<sup>3</sup> MUSSER	05	TWST	+ Surface $\mu^+$
0.72 ± 0.06 ± 0.08		AMORUSO	04	ICAR	Liquid Ar TPC
0.762 ± 0.008	170k	<sup>4</sup> FRYBERGER	68	ASPK	+ 25–53 MeV $e^+$
0.760 ± 0.009	280k	<sup>4</sup> SHERWOOD	67	ASPK	+ 25–53 MeV $e^+$
0.7503 ± 0.0026	800k	<sup>4</sup> PEOPLES	66	ASPK	+ 20–53 MeV $e^+$

<sup>1</sup> The quoted systematic error includes a contribution of 0.00013 (added in quadrature) from uncertainties on radiative corrections and on the Michel parameter  $\eta$ .

<sup>2</sup> The quoted systematic error includes a contribution of 0.00011 (added in quadrature) from the dependence on the Michel parameter  $\eta$ .

<sup>3</sup> The quoted systematic error includes a contribution of 0.00023 (added in quadrature) from the dependence on the Michel parameter  $\eta$ .

<sup>4</sup>  $\eta$  constrained = 0. These values incorporated into a two parameter fit to  $\rho$  and  $\eta$  by DERENZO 69.

### $\eta$ PARAMETER

(V–A) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.057 ± 0.034 OUR AVERAGE</b>					
0.071 ± 0.037 ± 0.005	30M	DANNEBERG	05	CNTR	+ 7–53 MeV $e^+$
0.011 ± 0.081 ± 0.026	5.3M	<sup>1</sup> BURKARD	85B	CNTR	+ 9–53 MeV $e^+$
–0.12 ± 0.21	6346	DERENZO	69	HBC	+ 1.6–6.8 MeV $e^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
–0.0021 ± 0.0070 ± 0.0010	30M	<sup>2</sup> DANNEBERG	05	CNTR	+ 7–53 MeV $e^+$
–0.012 ± 0.015 ± 0.003	5.3M	<sup>2</sup> BURKARD	85B	CNTR	+ 9–53 MeV $e^+$
–0.007 ± 0.013	5.3M	<sup>3</sup> BURKARD	85B	FIT	+ 9–53 MeV $e^+$
–0.7 ± 0.5	170k	<sup>4</sup> FRYBERGER	68	ASPK	+ 25–53 MeV $e^+$
–0.7 ± 0.6	280k	<sup>4</sup> SHERWOOD	67	ASPK	+ 25–53 MeV $e^+$
0.05 ± 0.5	800k	<sup>4</sup> PEOPLES	66	ASPK	+ 20–53 MeV $e^+$
–2.0 ± 0.9	9213	<sup>5</sup> PLANO	60	HBC	+ Whole spectrum

<sup>1</sup> Previously we used the global fit result from BURKARD 85B in OUR AVERAGE, we now only include their actual measurement.

<sup>2</sup>  $\alpha = \alpha' = 0$  assumed.

<sup>3</sup> Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

<sup>4</sup>  $\rho$  constrained = 0.75.

<sup>5</sup> Two parameter fit to  $\rho$  and  $\eta$ ; PLANO 60 discounts value for  $\eta$ .

### $\delta$ PARAMETER

(V–A) theory predicts  $\delta = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.75047 ± 0.00034 OUR AVERAGE</b>					
0.75049 ± 0.00021 ± 0.00027		<sup>1</sup> BAYES	11	TWST	+ Surface $\mu^+$
0.7486 ± 0.0026 ± 0.0028		<sup>2</sup> BALKE	88	SPEC	+ Surface $\mu^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.75067 ± 0.00030 ± 0.00067		MACDONALD	08	TWST	+ Surface $\mu^+$
0.74964 ± 0.00066 ± 0.00112	6G	<sup>3</sup> GAPONENKO	05	TWST	+ Surface $\mu^+$
		VOSSLER	69		
0.752 ± 0.009	490k	FRYBERGER	68	ASPK	+ 25–53 MeV $e^+$
0.782 ± 0.031		KRUGER	61		
0.78 ± 0.05	8354	PLANO	60	HBC	+ Whole spectrum

<sup>1</sup> The quoted systematic error includes a contribution of 0.00006 (added in quadrature) from uncertainties on radiative corrections and on the Michel parameter  $\eta$ .

<sup>2</sup> BALKE 88 uses  $\rho = 0.752 \pm 0.003$ .

<sup>3</sup> VOSSLER 69 has measured the asymmetry below 10 MeV. See comments about radiative corrections in VOSSLER 69.

### $|\xi|$ (PARAMETER) $\times (\mu$ LONGITUDINAL POLARIZATION)

(V–A) theory predicts  $\xi = 1$ , longitudinal polarization = 1.

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.0009 ± 0.0016 OUR AVERAGE</b>				
1.00084 ± 0.00029 ± 0.00165	BUENO	11	TWST	Surface $\mu^+$ beam
1.0027 ± 0.0079 ± 0.0030	BELTRAMI	87	CNTR	SIN, $\pi$ decay in flight
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.0003 ± 0.0006 ± 0.0038	JAMIESON	06	TWST	+ surface $\mu^+$ beam
1.0013 ± 0.0030 ± 0.0053	<sup>1</sup> IMAZATO	92	SPEC	+ $K^+ \rightarrow \mu^+ \nu_\mu$
0.975 ± 0.015	AKHMANOV	68	EMUL	140 kG
0.975 ± 0.030	GUREVICH	64	EMUL	See AKHMANOV 68
0.903 ± 0.027	<sup>2</sup> ALI-ZADE	61	EMUL	+ 27 kG
0.93 ± 0.06	PLANO	60	HBC	+ 8.8 kG
0.97 ± 0.05	BARDON	59	CNTR	Bromoform target

<sup>1</sup> The corresponding 90% confidence limit from IMAZATO 92 is  $|\xi P_\mu| > 0.990$ . This measurement is of  $K^+$  decay, not  $\pi^+$  decay, so we do not include it in an average, nor do we yet set up a separate data block for  $K$  results.

<sup>2</sup> Depolarization by medium not known sufficiently well.

$\xi \times (\mu \text{ LONGITUDINAL POLARIZATION}) \times \delta / \rho$ 

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.00179<math>\pm</math>0.00156 -0.00071</b>		<sup>1</sup> BAYES	11	TWST +	Surface $\mu^+$ beam
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>0.99682	90	<sup>2</sup> JODIDIO	86	SPEC +	TRIUMF
>0.9966	90	<sup>3</sup> STOKER	85	SPEC +	$\mu$ -spin rotation
>0.9959	90	CARR	83	SPEC +	11 kG

<sup>1</sup> BAYES 11 obtains the limit  $> 0.99909$  (90% CL) with the constraint that  $\xi \times (\mu \text{ LONGITUDINAL POLARIZATION}) \times \delta / \rho \leq 1.0$ .

<sup>2</sup> JODIDIO 86 includes data from CARR 83 and STOKER 85. The value here is from the erratum.

<sup>3</sup> STOKER 85 find  $(\xi P_{\mu} \delta / \rho) > 0.9955$  and  $> 0.9966$ , where the first limit is from new  $\mu$  spin-rotation data and the second is from combination with CARR 83 data. In  $V-A$  theory,  $(\delta / \rho) = 1.0$ .

 $\xi' = \text{LONGITUDINAL POLARIZATION OF } e^+$ 

( $V-A$ ) theory predicts the longitudinal polarization =  $\pm 1$  for  $e^{\pm}$ , respectively. We have flipped the sign for  $e^-$  so our programs can average.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.00 <math>\pm</math> 0.04 OUR AVERAGE</b>					
0.998 $\pm$ 0.045	1M	BURKARD	85	CNTR +	Bhabha + annihl
0.89 $\pm$ 0.28	29k	SCHWARTZ	67	OSPK -	Moller scattering
0.94 $\pm$ 0.38		BLOOM	64	CNTR +	Brems. transmiss.
1.04 $\pm$ 0.18		DUCLOS	64	CNTR +	Bhabha scattering
1.05 $\pm$ 0.30		BUHLER	63	CNTR +	Annihilation

 $\xi''$  PARAMETER

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.98 <math>\pm</math> 0.04 OUR AVERAGE</b>					
0.981 $\pm$ 0.045 $\pm$ 0.003	3.87M	PRIEELS	14	CNTR +	Bhabha + annihl
0.65 $\pm$ 0.36	326k	<sup>1</sup> BURKARD	85	CNTR +	Bhabha + annihl

<sup>1</sup> BURKARD 85 measure  $(\xi'' - \xi \xi') / \xi$  and  $\xi'$  and set  $\xi = 1$ .

TRANSVERSE  $e^+$  POLARIZATION IN PLANE OF  $\mu$  SPIN,  $e^+$  MOMENTUM

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>7 <math>\pm</math> 8 OUR AVERAGE</b>					
6.3 $\pm$ 7.7 $\pm$ 3.4	30M	DANNEBERG	05	CNTR +	7-53 MeV $e^+$
16 $\pm$ 21 $\pm$ 10	5.3M	BURKARD	85B	CNTR +	Annihil 9-53 MeV

TRANSVERSE  $e^+$  POLARIZATION NORMAL TO PLANE OF  $\mu$  SPIN,  $e^+$  MOMENTUM

Zero if  $T$  invariance holds.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-2 <math>\pm</math> 8 OUR AVERAGE</b>					
-3.7 $\pm$ 7.7 $\pm$ 3.4	30M	DANNEBERG	05	CNTR +	7-53 MeV $e^+$
7 $\pm$ 22 $\pm$ 7	5.3M	BURKARD	85B	CNTR +	Annihil 9-53 MeV

 $\alpha/A$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.4 <math>\pm</math> 4.3</b>		<sup>1</sup> BURKARD	85B	FIT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
15 $\pm$ 50 $\pm$ 14	5.3M	BURKARD	85B	CNTR +	9-53 MeV $e^+$

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 $\alpha'/A$ 

Zero if  $T$  invariance holds.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-10 <math>\pm</math> 20 OUR AVERAGE</b>					
-3.4 $\pm$ 21.3 $\pm$ 4.9	30M	DANNEBERG	05	CNTR +	7-53 MeV $e^+$
-47 $\pm$ 50 $\pm$ 14	5.3M	<sup>1</sup> BURKARD	85B	CNTR +	9-53 MeV $e^+$

• • • We do not use the following data for averages, fits, limits, etc. • • •

-0.2  $\pm$  4.3 <sup>2</sup> BURKARD 85B FIT

<sup>1</sup> Previously we used the global fit result from BURKARD 85B in OUR AVERAGE, we now only include their actual measurement. BURKARD 85B measure  $e^+$  polarizations  $P_{T_1}$  and  $P_{T_2}$  versus  $e^+$  energy.

<sup>2</sup> Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

 $\beta/A$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>3.9 <math>\pm</math> 6.2</b>		<sup>1</sup> BURKARD	85B	FIT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2 $\pm$ 17 $\pm$ 6	5.3M	BURKARD	85B	CNTR +	9-53 MeV $e^+$

<sup>1</sup> Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

 $\beta'/A$ 

Zero if  $T$  invariance holds.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2 <math>\pm</math> 7 OUR AVERAGE</b>					
-0.5 $\pm$ 7.8 $\pm$ 1.8	30M	DANNEBERG	05	CNTR +	7-53 MeV $e^+$
17 $\pm$ 17 $\pm$ 6	5.3M	<sup>1</sup> BURKARD	85B	CNTR +	9-53 MeV $e^+$

• • • We do not use the following data for averages, fits, limits, etc. • • •

-1.3  $\pm$  3.5  $\pm$  0.6 30M <sup>2</sup> DANNEBERG 05 CNTR + 7-53 MeV  $e^+$

1.5  $\pm$  6.3 <sup>3</sup> BURKARD 85B FIT

<sup>1</sup> Previously we used the global fit result from BURKARD 85B in OUR AVERAGE, we now only include their actual measurement. BURKARD 85B measure  $e^+$  polarizations  $P_{T_1}$  and  $P_{T_2}$  versus  $e^+$  energy.

<sup>2</sup>  $\alpha = \alpha' = 0$  assumed.

<sup>3</sup> Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

 $a/A$ 

This comes from an alternative parameterization to that used in the Summary Table (see the "Note on Muon Decay Parameters" above).

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<15.9	90	<sup>1</sup> BURKARD	85B FIT

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 $a'/A$ 

This comes from an alternative parameterization to that used in the Summary Table (see the "Note on Muon Decay Parameters" above).

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		
5.3 $\pm$ 4.1	<sup>1</sup> BURKARD	85B FIT

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 $(b'+b)/A$ 

This comes from an alternative parameterization to that used in the Summary Table (see the "Note on Muon Decay Parameters" above).

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<1.04	90	<sup>1</sup> BURKARD	85B FIT

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 $c/A$ 

This comes from an alternative parameterization to that used in the Summary Table (see the "Note on Muon Decay Parameters" above).

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<6.4	90	<sup>1</sup> BURKARD	85B FIT

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 $c'/A$ 

This comes from an alternative parameterization to that used in the Summary Table (see the "Note on Muon Decay Parameters" above).

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		
3.5 $\pm$ 2.0	<sup>1</sup> BURKARD	85B FIT

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 $\eta$  PARAMETER

( $V-A$ ) theory predicts  $\eta = 0$ .  $\eta$  affects spectrum of radiative muon decay.

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.02 <math>\pm</math> 0.08 OUR AVERAGE</b>				
-0.014 $\pm$ 0.090	EICHENBER...	84	ELEC +	$\rho$ free
+0.09 $\pm$ 0.14	BOGART	67	CNTR +	

• • • We do not use the following data for averages, fits, limits, etc. • • •

-0.035  $\pm$  0.098 EICHENBER... 84 ELEC +  $\rho=0.75$  assumed

 $\mu$  REFERENCES

MOHR	16	arXiv:1507.07956	P.J. Mohr, D.B. Newell, B.N. Taylor	(NIST)
Accepted for publication in RMP				
PRIEELS	14	PR D90 112003	R. Prieeels <i>et al.</i>	(LOUV, ETH, PSI+)
ADAM	13B	PRL 110 201801	J. Adam <i>et al.</i>	(MEG Collab.)
TISHCHENKO	13	PR D87 052003	V. Tishchenko <i>et al.</i>	(MuLan Collab.)
MOHR	12	RMP 84 1527	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
ADAM	11	PRL 107 171801	J. Adam <i>et al.</i>	(MEG Collab.)
BAYES	11	PRL 106 041804	R. Bayes <i>et al.</i>	(TWIST Collab.)
Also		PR D85 092013	A. Hillairet <i>et al.</i>	(TWIST Collab.)
BUENO	11	PR D84 032005	J.F. Bueno <i>et al.</i>	(TWIST Collab.)
Also		PR D85 039908 (err.)	J.F. Bueno <i>et al.</i>	(TWIST Collab.)
WEBBER	11	PRL 106 041803	D.M. Webber <i>et al.</i>	(MuLan Collab.)
Also		PRL 106 079901 (err.)	D.M. Webber <i>et al.</i>	(MuLan Collab.)
ADAM	10	NP B834 1	J. Adam <i>et al.</i>	(MEG Collab.)
BENNETT	09	PR D80 052008	G.W. Bennett <i>et al.</i>	(MUG-2 Collab.)
BARCZYK	08	PL B663 172	A. Barczyk <i>et al.</i>	(FAST Collab.)
MACDONALD	08	PR D78 032010	R.P. MacDonald <i>et al.</i>	(TWIST Collab.)

# Lepton Particle Listings

$\mu, \tau$

MOHR	08	RMP 80 633	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
CHITWOOD	07	PRL 99 032001	D.B. Chitwood <i>et al.</i>	(MULAN Collab.)
BENNETT	06	PR D73 072003	G.W. Bennett <i>et al.</i>	(MUG-2 Collab.)
BERTL	06	EPJ C47 337	W. Bertl <i>et al.</i>	(SINDRUM II Collab.)
JAMIESON	06	PR D74 072007	B. Jamieson <i>et al.</i>	(TWIST Collab.)
DANNEBERG	05	PRL 94 021802	N. Danneberg <i>et al.</i>	(ETH, JAGL, PSI+)
GAPONENKO	05	PR D71 071101	A. Gaponenko <i>et al.</i>	(TWIST Collab.)
MOHR	05	RMP 77 1	P.J. Mohr, B.N. Taylor	(NIST)
MUSSER	05	PRL 94 101805	J.R. Musser <i>et al.</i>	(TWIST Collab.)
AMORUSO	04	EPJ C33 233	S. Amoruso <i>et al.</i>	(ICARUS Collab.)
BENNETT	04	PRL 92 161802	G.W. Bennett <i>et al.</i>	(Muon(g-2) Collab.)
AHMED	02	PR D65 112002	M. Ahmed <i>et al.</i>	(MEGA Collab.)
BENNETT	02	PRL 89 101804	G.W. Bennett <i>et al.</i>	(Muon(g-2) Collab.)
BROWN	01	PRL 86 2227	H.N. Brown <i>et al.</i>	(Muon(g-2) Collab.)
BROWN	00	PR D62 091101	H.N. Brown <i>et al.</i>	(BNL/G-2 Collab.)
MEYER	00	PRL 84 1136	V. Meyer <i>et al.</i>	
BROOKS	99	PRL 83 1521	M.L. Brooks <i>et al.</i>	(MEGA/LAMPF Collab.)
HUGHES	99	RMP 71 5133	V.W. Hughes, T. Kinoshita	
LIU	99	PRL 82 711	W. Liu <i>et al.</i>	(LAMPF Collab.)
MOHR	99	JPCRD 28 1713	P.J. Mohr, B.N. Taylor	(NIST)
Also		RMP 72 351	P.J. Mohr, B.N. Taylor	(NIST)
WILLMANN	99	PRL 82 49	L. Willmann <i>et al.</i>	
FELDMAN	98	PR D57 3873	G.J. Feldman, R.D. Cousins	
KAULARD	98	PL B422 334	J. Kaulard <i>et al.</i>	(SINDRUM-II Collab.)
GORDEEV	97	PAN 60 1164	V.A. Gordeev <i>et al.</i>	(PNPI)
Also		Translated from YAF 60 1291.		
ABELA	96	PRL 77 1950	R. Abela <i>et al.</i>	(PSI, ZURI, HEIDH, TBIL+)
HONECKER	96	PRL 76 200	W. Honecker <i>et al.</i>	(SINDRUM II Collab.)
DOHMEN	93	PL B317 631	C. Dohmen <i>et al.</i>	(PSI SINDRUM II Collab.)
FREEDMAN	93	PR D47 811	S.J. Freedman <i>et al.</i>	(LAMPF E645 Collab.)
NI	93	PR D48 1976	B. Ni <i>et al.</i>	(LAMPF Crystal-Box Collab.)
IMAZATO	92	PRL 69 877	J. Imaizoto <i>et al.</i>	(KEK, INUS, TOKY+)
BARANOV	91	SJNP 53 802	V.A. Baranov <i>et al.</i>	(JINR)
Also		Translated from YAF 53 1302.		
KRAKAUER	91B	PL B263 534	D.A. Krakauer <i>et al.</i>	(UMD, UCI, LANL)
MATTHIAS	91	PRL 66 2716	B.E. Matthias <i>et al.</i>	(YALE, HEIDP, WILL+)
Also		PRL 67 932 (erratum)	B.E. Matthias <i>et al.</i>	(YALE, HEIDP, WILL+)
HUBER	90B	PR D41 2709	T.M. Huber <i>et al.</i>	(WYOM, VICT, ARIZ+)
AHMAD	88	PR D38 2102	S. Ahmad <i>et al.</i>	(TRIUM, VICT, VPL, BRCO+)
Also		PRL 59 970	S. Ahmad <i>et al.</i>	(TRIUM, VPI, VICT, BRCO+)
BALKE	88	PR D37 587	B. Balke <i>et al.</i>	(LBL, UCB, COLO, NWES+)
BELLGARDT	88	NP B299 1	U. Bellgardt <i>et al.</i>	(SINDRUM Collab.)
BOLTON	88	PR D38 2077	R.D. Bolton <i>et al.</i>	(LANL, STAN, CHIC+)
Also		PRL 56 2461	R.D. Bolton <i>et al.</i>	(LANL, STAN, CHIC+)
Also		PRL 57 3241	D. Grosnick <i>et al.</i>	(CHIC, LANL, STAN+)
BELTRAMI	87	PL B194 326	I. Beltrami <i>et al.</i>	(ETH, SIN, MANZ)
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RSC, NBS)
BEER	86	PRL 57 671	G.A. Beer <i>et al.</i>	(VICT, TRIUM, WYOM)
JODIDIO	86	PR D34 1967	A. Jodidio <i>et al.</i>	(A. Jodidio, NWES, TRIUM)
Also		PR D37 237 (erratum)	A. Jodidio <i>et al.</i>	(SINDRUM Collab.)
BERTL	85	NP B260 1	W. Bertl <i>et al.</i>	(SINDRUM Collab.)
BRYMAN	85	PRL 55 465	D.A. Bryman <i>et al.</i>	(TRIUM, CNRC, BRCO+)
BURKARD	85	PL 150B 242	H. Burkhardt <i>et al.</i>	(ETH, SIN, MANZ)
BURKARD	85B	PL 160B 343	H. Burkhardt <i>et al.</i>	(ETH, SIN, MANZ)
Also		PR D24 2004	F. Corribeau <i>et al.</i>	(ETH, SIN, MANZ)
Also		PL 129B 260	F. Corribeau <i>et al.</i>	(ETH, SIN, MANZ)
STOKER	85	PRL 54 1887	D.P. Stoker <i>et al.</i>	(LBL, NWES, TRIUM)
BARDIN	84	PL 137B 135	G. Bardin <i>et al.</i>	(SACL, CERN, BGNA, FIRZ)
BERTL	84	PL 140B 299	W. Bertl <i>et al.</i>	(SINDRUM Collab.)
BOLTON	84	PRL 53 1415	R.D. Bolton <i>et al.</i>	(LANL, CHIC, STAN+)
EICHENBER...	84	NP A412 523	W. Eichenberger, R. Engfer, A. van der Schaff	
GIOVANETTI	84	PR D29 343	K.L. Giovanetti <i>et al.</i>	(WILL)
AZUELOS	83	PRL 51 164	G. Azuelos <i>et al.</i>	(MONT, TRIUM, BRCO)
Also		PRL 39 1113	P. Depommier <i>et al.</i>	(MONT, BRCO, TRIUM+)
BERGSM	83	PL 122B 465	F. Bergsma <i>et al.</i>	(CHARM Collab.)
CARR	83	PRL 51 627	J. Carr <i>et al.</i>	(LBL, NWES, TRIUM)
KINNISON	82	PR D25 2846	W.W. Kinnison <i>et al.</i>	(EFI, STAN, LANL)
Also		PRL 42 556	J.D. Bowman <i>et al.</i>	(LASL, EFI, STAN)
KLEMP	82	PR D25 652	E. Klemp <i>et al.</i>	(MANZ, ETH)
MARIAM	82	PRL 49 993	F.G. Mariam <i>et al.</i>	(YALE, HEIDH, BERN)
MARSHALL	82	PR D25 1174	G.M. Marshall <i>et al.</i>	(BRCO)
NEMETHY	81	CNPP 10 147	P. Nemethy, V.W. Hughes	(LBL, YALE)
ABELA	80	PL 95B 318	R. Abela <i>et al.</i>	(BASL, KARLK, KARLE)
BADERT...	80	LNC 28 401	A. Badertscher <i>et al.</i>	(BERN)
Also		NP A377 406	A. Badertscher <i>et al.</i>	(BERN)
JONKER	80	PL 93B 203	M. Jonker <i>et al.</i>	(CHARM Collab.)
SCHAAF	80	NP A340 249	A. van der Schaaf <i>et al.</i>	(ZURI, ETH+)
Also		PL 72B 183	H.P. Povel <i>et al.</i>	(ZURI, ETH, SIN)
WILLIS	80	PRL 44 522	S.E. Willis <i>et al.</i>	(YALE, LBL, LASL+)
Also		PRL 45 1370	S.E. Willis <i>et al.</i>	(YALE, LBL, LASL+)
BAILEY	79	NP B150 1	J.M. Bailey	(CERN, DARE, MANZ)
BADERT...	78	PL 79B 371	A. Badertscher <i>et al.</i>	(BERN)
BAILEY	78	IP 64 345	J.M. Bailey	(DARE, BERN, SHEF, MANZ, RMCs+)
Also		NP B150 1	J.M. Bailey	(CERN, DARE, MANZ)
BLIETSCHAU	78	NP B133 205	J. Blietschau <i>et al.</i>	(Gargamelle Collab.)
BOWMAN	78	PRL 41 442	J.D. Bowman <i>et al.</i>	(LASL, IAS, CMU+)
CAMANI	78	PL 77B 326	M. Camani <i>et al.</i>	(ETH, MANZ)
BADERT...	77	PRL 39 1385	A. Badertscher <i>et al.</i>	(BERN)
CASPERSON	77	PRL 38 956	D.E. Casperson <i>et al.</i>	(BERN, HEIDH, LASL+)
DEPOMMIER	77	PRL 39 1113	P. Depommier <i>et al.</i>	(MONT, BRCO, TRIUM+)
BALANDIN	74	JETP 40 811	M.P. Balandin <i>et al.</i>	(JINR)
Also		Translated from ZETF 67 1631.		
COHEN	73	JPCRD 2 664	E.R. Cohen, B.N. Taylor	(RSC, NBS)
DUCLOS	73	PL 47B 491	J. Duclos, A. Magnon, J. Picard	(SACL)
EICHEN	73	PL 46B 281	T. Eichten <i>et al.</i>	(Gargamelle Collab.)
BRYMAN	72	PRL 28 1469	D.A. Bryman <i>et al.</i>	(VPI)
CROWE	72	PR D5 2145	K.M. Crowe <i>et al.</i>	(LBL, WASH)
CRANE	71	PRL 27 474	T. Crane <i>et al.</i>	(YALE)
DERENZO	69	PR 181 1854	S.E. Derenzo	(EFI)
VOSSLER	69	NC 63A 423	C. Vossler	(EFI)
AKHMANOV	68	SJNP 6 230	V.V. Akhmanov <i>et al.</i>	(KIAE)
Also		Translated from YAF 6 316.		
FRYBERGER	68	PR 166 1379	D. Fryberger	(EFI)
BOGART	67	PR 156 1405	E. Bogart <i>et al.</i>	(COLU)
SCHWARTZ	67	PR 162 1306	D.M. Schwartz	(EFI)
SHERWOOD	67	PR 156 1475	B.A. Sherwood	(EFI)
PEOPLES	66	Nevis 147 unpub.	J. Peoples	(COLU)
BLOOM	64	PL 8 87	S. Bloom <i>et al.</i>	(CERN)
DUCLOS	64	PL 9 62	J. Duclos <i>et al.</i>	(CERN)
GUREVICH	64	PL 11 185	I.I. Gurevich <i>et al.</i>	(KIAE)
BUHLER	63	PL 7 368	A. Buhler-Broglin <i>et al.</i>	(CERN)
MEYER	63	PR 132 2693	S.L. Meyer <i>et al.</i>	(COLU)
CHARPAK	62	PL 1 16	G. Charpak <i>et al.</i>	(CERN)
CONFORTO	62	NC 26 261	G. Conforto <i>et al.</i>	(INFN, ROMA, CERN)
ALI-ZADE	61	JETP 33 313	S.A. Ali-Zade, I.I. Gurevich, B.A. Nikolsky	
Also		Translated from ZETF 40 452.		
CRITTENDEN	61	PR 121 1823	R.R. Crittenden, W.D. Walker, J. Ballam	(WISC+)
KRUGER	61	UCRL 9322 unpub.	H. Kruger	(LRL)
GUREVICH	60	JETP 10 225	I.I. Gurevich, B.A. Nikolsky, L.V. Surkova	(ITEP)
Also		Translated from ZETF 37 318.		

PLANO	60	PR 119 1400	R.J. Plano	(COLU)
ASHKIN	59	NC 14 1266	J. Ashkin <i>et al.</i>	(CERN)
BARDON	59	PRL 2 56	M. Bardon, D. Berley, L.M. Lederman	(COLU)
LEE	59	PRL 3 55	J. Lee, N.P. Samios	(COLU)

$\tau$

$$J = \frac{1}{2}$$

$\tau$  discovery paper was PERL 75.  $e^+e^- \rightarrow \tau^+\tau^-$  cross-section threshold behavior and magnitude are consistent with pointlike spin-1/2 Dirac particle. BRANDELIK 78 ruled out pointlike spin-0 or spin-1 particle. FELDMAN 78 ruled out  $J = 3/2$ . KIRKBY 79 also ruled out  $J$ =integer,  $J = 3/2$ .

## $\tau$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1776.86 ± 0.12 OUR AVERAGE</b>				
1776.91 ± 0.12 $^{+0.10}_{-0.13}$	1171	<sup>1</sup> ABLIKIM	14D BES3	23.3 pb <sup>-1</sup> , $E_{cm}^{ee} = 3.54\text{--}3.60$ GeV
1776.68 ± 0.12 ± 0.41	682k	<sup>2</sup> AUBERT	09AK BABR	423 fb <sup>-1</sup> , $E_{cm}^{ee} = 10.6$ GeV
1776.81 ± 0.25 $^{+0.25}_{-0.23}$ ± 0.15	81	ANASHIN	07 KEDR	6.7 pb <sup>-1</sup> , $E_{cm}^{ee} = 3.54\text{--}3.78$ GeV
1776.61 ± 0.13 ± 0.35		<sup>2</sup> BELOUS	07 BELL	414 fb <sup>-1</sup> , $E_{cm}^{ee} = 10.6$ GeV
1775.1 ± 1.6 ± 1.0	13.3k	<sup>3</sup> ABBIENDI	00A OPAL	1990–1995 LEP runs
1778.2 ± 0.8 ± 1.2		ANASTASSOV	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV
1776.96 ± 0.18 ± 0.25 $^{+0.25}_{-0.21}$ ± 0.17	65	<sup>4</sup> BAI	96 BES	$E_{cm}^{ee} = 3.54\text{--}3.57$ GeV
1776.3 ± 2.4 ± 1.4	11k	<sup>5</sup> ALBRECHT	92M ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
1783 $^{+3}_{-4}$	692	<sup>6</sup> BACINO	78B DLCO	$E_{cm}^{ee} = 3.1\text{--}7.4$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1777.8 ± 0.7 ± 1.7	35k	<sup>7</sup> BALEST	93 CLEO	Repl. by ANASTASSOV 97
1776.9 $^{+0.4}_{-0.5}$ ± 0.2	14	<sup>8</sup> BAI	92 BES	Repl. by BAI 96

- <sup>1</sup> ABLIKIM 14D fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  at different energies near threshold.
- <sup>2</sup> AUBERT 09AK and BELOUS 07 fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi^+\pi^-\nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .
- <sup>3</sup> ABBIENDI 00A fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi^\pm \leq 2\pi^0 \nu_\tau$  and  $\tau \rightarrow \pi^\pm \pi^+\pi^- \leq 1\pi^0 \nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .
- <sup>4</sup> BAI 96 fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  at different energies near threshold.
- <sup>5</sup> ALBRECHT 92M fit  $\tau$  pseudomass spectrum in  $\tau^- \rightarrow 2\pi^-\pi^+\nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .
- <sup>6</sup> BACINO 78B value comes from  $e^\pm X^\mp$  threshold. Published mass 1782 MeV increased by 1 MeV using the high precision  $\psi(2S)$  mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.
- <sup>7</sup> BALEST 93 fit spectra of minimum kinematically allowed  $\tau$  mass in events of the type  $e^+e^- \rightarrow \tau^+\tau^- \rightarrow (\pi^+ n\pi^0 \nu_\tau)(\pi^- m\pi^0 \nu_\tau)$   $n \leq 2, m \leq 2, 1 \leq n+m \leq 3$ . If  $m_{\nu_\tau} \neq 0$ , result increases by  $(m_{\nu_\tau}^2/1100 \text{ MeV})$ .
- <sup>8</sup> BAI 92 fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  near threshold using  $e\mu$  events.

## $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}}$

A test of  $CPT$  invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.8 × 10<sup>-4</sup></b>	90	BELOUS	07 BELL	414 fb <sup>-1</sup> , $E_{cm}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.5 × 10 <sup>-4</sup>	90	<sup>1</sup> AUBERT	09AK BABR	423 fb <sup>-1</sup> , $E_{cm}^{ee} = 10.6$ GeV
<3.0 × 10 <sup>-3</sup>	90	ABBIENDI	00A OPAL	1990–1995 LEP runs
<sup>1</sup> AUBERT 09AK quote both the listed upper limit and $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}} = (-3.4 \pm 1.3 \pm 0.3) \times 10^{-4}$ .				

## $\tau$ MEAN LIFE

VALUE (10 <sup>-15</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>290.3 ± 0.5 OUR AVERAGE</b>				
290.17 ± 0.53 ± 0.33 1.1M		BELOUS	14 BELL	711 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
290.9 ± 1.4 ± 1.0		ABDALLAH	04T DLPH	1991–1995 LEP runs
293.2 ± 2.0 ± 1.5		ACCIARRI	00B L3	1991–1995 LEP runs
290.1 ± 1.5 ± 1.1		BARATE	97R ALEP	1989–1994 LEP runs
289.2 ± 1.7 ± 1.2		ALEXANDER	96E OPAL	1990–1994 LEP runs
289.0 ± 2.8 ± 4.0 57.4k		BALEST	96 CLEO	$E_{cm}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
291.2 ± 2.0 ± 1.2		BARATE	97I ALEP	Repl. by BARATE 97R
291.4 ± 3.0		ABREU	96B DLPH	Repl. by ABDALLAH 04T
290.1 ± 4.0	34k	ACCIARRI	96k L3	Repl. by ACCIARRI 00B
297 ± 9 ± 5	1671	ABE	95Y SLD	1992–1993 SLC runs
304 ± 14 ± 7	4100	BATTLE	92 CLEO	$E_{cm}^{ee} = 10.6$ GeV
301 ± 29 ± 3780		KLEINWORT	89 JADE	$E_{cm}^{ee} = 35\text{--}46$ GeV
288 ± 16 ± 17	807	AMIDEI	88 MRK2	$E_{cm}^{ee} = 29$ GeV

306	±20	±14	695	BRAUNSCH...	88c	TASS	$E_{\text{cm}}^{\text{ee}} = 36 \text{ GeV}$
299	±15	±10	1311	ABACHI	87c	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
295	±14	±11	5696	ALBRECHT	87p	ARG	$E_{\text{cm}}^{\text{ee}} = 9.3\text{--}10.6 \text{ GeV}$
309	±17	±7	3788	BAND	87b	MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
325	±14	±18	8470	BEBEK	87c	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
460 ± 190			102	FELDMAN	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

$$(\tau_{\tau^+} - \tau_{\tau^-}) / \tau_{\text{average}}$$

Test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.0 × 10<sup>-3</sup></b>	90	<sup>1</sup> BELOUS	14	BELL $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup>BELOUS 14 quote limit on the absolute value of the relative lifetime difference.

### $\tau$ MAGNETIC MOMENT ANOMALY

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

$$\mu_{\tau}/(e\hbar/2m_{\tau})-1 = (g_{\tau}-2)/2$$

For a theoretical calculation  $[(g_{\tau}-2)/2 = 117\,721(5) \times 10^{-8}]$ , see EIDELMAN 07.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt; -0.052 and &lt; 0.013 (CL = 95%) OUR LIMIT</b>				
> -0.052 and < 0.013	95	<sup>1</sup> ABDALLAH	04k	DLPH $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.107	95	<sup>2</sup> ACHARD	04g	L3 $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
> -0.007 and < 0.005	95	<sup>3</sup> GONZALEZ-S..00	RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ and $W \rightarrow \tau\nu_{\tau}$
> -0.052 and < 0.058	95	<sup>4</sup> ACCIARRI	98e	L3 1991-1995 LEP runs
> -0.068 and < 0.065	95	<sup>5</sup> ACKERSTAFF	98n	OPAL 1990-1995 LEP runs
> -0.004 and < 0.006	95	<sup>6</sup> ESCRIBANO	97	RVUE $Z \rightarrow \tau^+\tau^-$ at LEP
<0.01	95	<sup>7</sup> ESCRIBANO	93	RVUE $Z \rightarrow \tau^+\tau^-$ at LEP
<0.12	90	GRIFOLS	91	RVUE $Z \rightarrow \tau\tau\gamma$ at LEP
<0.023	95	<sup>8</sup> SILVERMAN	83	RVUE $e^+e^- \rightarrow \tau^+\tau^-$ at PETRA

<sup>1</sup>ABDALLAH 04k limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV. In addition to the limits, the authors also quote a value of  $-0.018 \pm 0.017$ .

<sup>2</sup>ACHARD 04g limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of the magnetic moment anomaly.

<sup>3</sup>GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

<sup>4</sup>ACCIARRI 98e use  $Z \rightarrow \tau^+\tau^-\gamma$  events. In addition to the limits, the authors also quote a value of  $0.004 \pm 0.027 \pm 0.023$ .

<sup>5</sup>ACKERSTAFF 98n use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_{\tau}^2$  to  $(M_Z - m_{\tau})^2$ .

<sup>6</sup>ESCRIBANO 97 use preliminary experimental results.

<sup>7</sup>ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+\tau^-)$ , and is on the absolute value of the magnetic moment anomaly.

<sup>8</sup>SILVERMAN 83 limit is derived from  $e^+e^- \rightarrow \tau^+\tau^-$  total cross-section measurements for  $q^2$  up to  $(37 \text{ GeV})^2$ .

### $\tau$ ELECTRIC DIPOLE MOMENT ( $d_{\tau}$ )

A nonzero value is forbidden by both *T* invariance and *P* invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

#### Re( $d_{\tau}$ )

VALUE (10 <sup>-16</sup> e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>- 0.22 to 0.45</b>	95	<sup>1</sup> INAMI	03	BELL $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 2.3	90	<sup>2</sup> GROZIN	09A	RVUE From $e$ EDM limit
< 3.7	95	<sup>3</sup> ABDALLAH	04k	DLPH $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
< 11.4	95	<sup>4</sup> ACHARD	04g	L3 $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
< 4.6	95	<sup>5</sup> ALBRECHT	00	ARG $E_{\text{cm}}^{\text{ee}} = 10.4 \text{ GeV}$
> -3.1 and < 3.1	95	ACCIARRI	98e	L3 1991-1995 LEP runs
> -3.8 and < 3.6	95	ACKERSTAFF	98n	OPAL 1990-1995 LEP runs
< 0.11	95	<sup>7,8</sup> ESCRIBANO	97	RVUE $Z \rightarrow \tau^+\tau^-$ at LEP
< 0.5	95	<sup>9</sup> ESCRIBANO	93	RVUE $Z \rightarrow \tau^+\tau^-$ at LEP
< 7	90	GRIFOLS	91	RVUE $Z \rightarrow \tau\tau\gamma$ at LEP
< 1.6	90	DELAGUILA	90	RVUE $e^+e^- \rightarrow \tau^+\tau^-$ $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$

<sup>1</sup>INAMI 03 use  $e^+e^- \rightarrow \tau^+\tau^-$  events.

<sup>2</sup>GROZIN 09A calculate the contribution to the electron electric dipole moment from the  $\tau$  electric dipole moment appearing in loops, which is  $\Delta d_e = 6.9 \times 10^{-12} d_{\tau}$ . Dividing the REGAN 02 upper limit  $|d_e| \leq 1.6 \times 10^{-27} \text{ e cm}$  at CL=90% by  $6.9 \times 10^{-12}$  gives this limit.

<sup>3</sup>ABDALLAH 04k limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV and is on the absolute value of  $d_{\tau}$ .

<sup>4</sup>ACHARD 04g limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of  $d_{\tau}$ .

<sup>5</sup>ALBRECHT 00 use  $e^+e^- \rightarrow \tau^+\tau^-$  events. Limit is on the absolute value of  $\text{Re}(d_{\tau})$ .

<sup>6</sup>ACKERSTAFF 98n use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_{\tau}^2$  to  $(M_Z - m_{\tau})^2$ .

<sup>7</sup>ESCRIBANO 97 derive the relationship  $|d_{\tau}| = \cot \theta_W |d_{\tau}^W|$  using effective Lagrangian methods, and use a conference result  $|d_{\tau}^W| < 5.8 \times 10^{-18} \text{ e cm}$  at 95% CL (L. Silvestris, ICHEP96) to obtain this result.

<sup>8</sup>ESCRIBANO 97 use preliminary experimental results.

<sup>9</sup>ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+\tau^-)$ , and is on the absolute value of the electric dipole moment.

#### Im( $d_{\tau}$ )

VALUE (10 <sup>-16</sup> e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>- 0.25 to 0.008</b>	95	<sup>1</sup> INAMI	03	BELL $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.8	95	<sup>2</sup> ALBRECHT	00	ARG $E_{\text{cm}}^{\text{ee}} = 10.4 \text{ GeV}$

<sup>1</sup>INAMI 03 use  $e^+e^- \rightarrow \tau^+\tau^-$  events.

<sup>2</sup>ALBRECHT 00 use  $e^+e^- \rightarrow \tau^+\tau^-$  events. Limit is on the absolute value of  $\text{Im}(d_{\tau})$ .

### $\tau$ WEAK DIPOLE MOMENT ( $d_{\tau}^W$ )

A nonzero value is forbidden by *CP* invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

#### Re( $d_{\tau}^W$ )

VALUE (10 <sup>-17</sup> e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.50</b>	95	<sup>1</sup> HEISTER	03f	ALEP 1990-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0	90	<sup>1</sup> ACCIARRI	98c	L3 1991-1995 LEP runs
<0.56	95	ACKERSTAFF	97L	OPAL 1991-1995 LEP runs
<0.78	95	<sup>2</sup> AKERS	95f	OPAL Repl. by ACKERSTAFF 97L
<1.5	95	<sup>2</sup> BUSKULIC	95c	ALEP Repl. by HEISTER 03f
<7.0	95	<sup>2</sup> ACTON	92f	OPAL $Z \rightarrow \tau^+\tau^-$ at LEP
<3.7	95	<sup>2</sup> BUSKULIC	92j	ALEP Repl. by BUSKULIC 95c

<sup>1</sup>Limit is on the absolute value of the real part of the weak dipole moment.

<sup>2</sup>Limit is on the absolute value of the real part of the weak dipole moment, and applies for  $q^2 = m_Z^2$ .

#### Im( $d_{\tau}^W$ )

VALUE (10 <sup>-17</sup> e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1</b>	95	<sup>1</sup> HEISTER	03f	ALEP 1990-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.5	95	ACKERSTAFF	97L	OPAL 1991-1995 LEP runs
<4.5	95	<sup>2</sup> AKERS	95f	OPAL Repl. by ACKERSTAFF 97L

<sup>1</sup>HEISTER 03f limit is on the absolute value of the imaginary part of the weak dipole moment.

<sup>2</sup>Limit is on the absolute value of the imaginary part of the weak dipole moment, and applies for  $q^2 = m_Z^2$ .

### $\tau$ WEAK ANOMALOUS MAGNETIC DIPOLE MOMENT ( $a_{\tau}^W$ )

Electroweak radiative corrections are expected to contribute at the  $10^{-6}$  level. See BERNABEU 95.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

#### Re( $a_{\tau}^W$ )

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 × 10<sup>-3</sup></b>	95	<sup>1</sup> HEISTER	03f	ALEP 1990-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> -0.0024 and < 0.0025	95	<sup>2</sup> GONZALEZ-S..00	RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ and $W \rightarrow \tau\nu_{\tau}$
<4.5 × 10 <sup>-3</sup>	90	<sup>1</sup> ACCIARRI	98c	L3 1991-1995 LEP runs

<sup>1</sup>Limit is on the absolute value of the real part of the weak anomalous magnetic dipole moment.

<sup>2</sup>GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

# Lepton Particle Listings

$\tau$

$\text{Im}(\alpha_\tau^W)$	CL%	DOCUMENT ID	TECN	COMMENT
$<2.7 \times 10^{-3}$	95	<sup>1</sup> HEISTER	03F ALEP	1990-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<9.9 \times 10^{-3}$	90	<sup>1</sup> ACCIARRI	98c L3	1991-1995 LEP runs
<sup>1</sup> Limit is on the absolute value of the imaginary part of the weak anomalous magnetic dipole moment.				

## $\tau^-$ DECAY MODES

$\tau^+$  modes are charge conjugates of the modes below. “ $h^\pm$ ” stands for  $\pi^\pm$  or  $K^\pm$ . “ $\ell$ ” stands for e or  $\mu$ . “Neutrals” stands for  $\gamma$ ’s and/or  $\pi^0$ ’s.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Modes with one charged particle</b>		
$\Gamma_1$ particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ (“1-prong”)	(85.24 $\pm$ 0.06) %	
$\Gamma_2$ particle $^- \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	(84.58 $\pm$ 0.06) %	
$\Gamma_3$ $\mu^- \overline{\nu}_\mu \nu_\tau$	[a] (17.39 $\pm$ 0.04) %	
$\Gamma_4$ $\mu^- \overline{\nu}_\mu \nu_\tau \gamma$	[b] ( 3.68 $\pm$ 0.10 ) $\times 10^{-3}$	
$\Gamma_5$ $e^- \overline{\nu}_e \nu_\tau$	[a] (17.82 $\pm$ 0.04) %	
$\Gamma_6$ $e^- \overline{\nu}_e \nu_\tau \gamma$	[b] ( 1.84 $\pm$ 0.05 ) %	
$\Gamma_7$ $h^- \geq 0 K_L^0 \nu_\tau$	(12.03 $\pm$ 0.05) %	
$\Gamma_8$ $h^- \nu_\tau$	(11.51 $\pm$ 0.05) %	
$\Gamma_9$ $\pi^- \nu_\tau$	[a] (10.82 $\pm$ 0.05) %	
$\Gamma_{10}$ $K^- \nu_\tau$	[a] ( 6.96 $\pm$ 0.10 ) $\times 10^{-3}$	
$\Gamma_{11}$ $h^- \geq 1$ neutrals $\nu_\tau$	(37.00 $\pm$ 0.09) %	
$\Gamma_{12}$ $h^- \geq 1 \pi^0 \nu_\tau$ (ex. $K^0$ )	(36.51 $\pm$ 0.09) %	
$\Gamma_{13}$ $h^- \pi^0 \nu_\tau$	(25.93 $\pm$ 0.09) %	
$\Gamma_{14}$ $\pi^- \pi^0 \nu_\tau$	[a] (25.49 $\pm$ 0.09) %	
$\Gamma_{15}$ $\pi^- \pi^0$ non- $\rho(770) \nu_\tau$	( 3.0 $\pm$ 3.2 ) $\times 10^{-3}$	
$\Gamma_{16}$ $K^- \pi^0 \nu_\tau$	[a] ( 4.33 $\pm$ 0.15 ) $\times 10^{-3}$	
$\Gamma_{17}$ $h^- \geq 2 \pi^0 \nu_\tau$	(10.81 $\pm$ 0.09) %	
$\Gamma_{18}$ $h^- 2 \pi^0 \nu_\tau$	( 9.48 $\pm$ 0.10 ) %	
$\Gamma_{19}$ $h^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 9.32 $\pm$ 0.10 ) %	
$\Gamma_{20}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 9.26 $\pm$ 0.10 ) %	
$\Gamma_{21}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), scalar	< 9 $\times 10^{-3}$	CL=95%
$\Gamma_{22}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), vector	< 7 $\times 10^{-3}$	CL=95%
$\Gamma_{23}$ $K^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 6.5 $\pm$ 2.2 ) $\times 10^{-4}$	
$\Gamma_{24}$ $h^- \geq 3 \pi^0 \nu_\tau$	( 1.34 $\pm$ 0.07 ) %	
$\Gamma_{25}$ $h^- \geq 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.25 $\pm$ 0.07 ) %	
$\Gamma_{26}$ $h^- 3 \pi^0 \nu_\tau$	( 1.18 $\pm$ 0.07 ) %	
$\Gamma_{27}$ $\pi^- 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 1.04 $\pm$ 0.07 ) %	
$\Gamma_{28}$ $K^- 3 \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	[a] ( 4.8 $\pm$ 2.1 ) $\times 10^{-4}$	
$\Gamma_{29}$ $h^- 4 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.6 $\pm$ 0.4 ) $\times 10^{-3}$	
$\Gamma_{30}$ $h^- 4 \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	[a] ( 1.1 $\pm$ 0.4 ) $\times 10^{-3}$	
$\Gamma_{31}$ $a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau$	( 3.8 $\pm$ 1.5 ) $\times 10^{-4}$	
$\Gamma_{32}$ $K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau$	( 1.552 $\pm$ 0.029) %	
$\Gamma_{33}$ $K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau$	( 8.59 $\pm$ 0.28 ) $\times 10^{-3}$	
<b>Modes with <math>K^0</math>’s</b>		
$\Gamma_{34}$ $K_S^0$ (particles) $^- \nu_\tau$	( 9.44 $\pm$ 0.28 ) $\times 10^{-3}$	
$\Gamma_{35}$ $h^- \overline{K}^0 \nu_\tau$	( 9.87 $\pm$ 0.14 ) $\times 10^{-3}$	
$\Gamma_{36}$ $\pi^- \overline{K}^0 \nu_\tau$	[a] ( 8.40 $\pm$ 0.14 ) $\times 10^{-3}$	
$\Gamma_{37}$ $\pi^- \overline{K}^0$ (non- $K^*(892)^- \nu_\tau$	( 5.4 $\pm$ 2.1 ) $\times 10^{-4}$	
$\Gamma_{38}$ $K^- K^0 \nu_\tau$	[a] ( 1.48 $\pm$ 0.05 ) $\times 10^{-3}$	
$\Gamma_{39}$ $K^- K^0 \geq 0 \pi^0 \nu_\tau$	( 2.98 $\pm$ 0.08 ) $\times 10^{-3}$	
$\Gamma_{40}$ $h^- \overline{K}^0 \pi^0 \nu_\tau$	( 5.32 $\pm$ 0.13 ) $\times 10^{-3}$	
$\Gamma_{41}$ $\pi^- \overline{K}^0 \pi^0 \nu_\tau$	[a] ( 3.82 $\pm$ 0.13 ) $\times 10^{-3}$	
$\Gamma_{42}$ $\overline{K}^0 \rho^- \nu_\tau$	( 2.2 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{43}$ $K^- K^0 \pi^0 \nu_\tau$	[a] ( 1.50 $\pm$ 0.07 ) $\times 10^{-3}$	
$\Gamma_{44}$ $\pi^- \overline{K}^0 \geq 1 \pi^0 \nu_\tau$	( 4.08 $\pm$ 0.25 ) $\times 10^{-3}$	
$\Gamma_{45}$ $\pi^- \overline{K}^0 \pi^0 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 2.6 $\pm$ 2.3 ) $\times 10^{-4}$	
$\Gamma_{46}$ $K^- K^0 \pi^0 \pi^0 \nu_\tau$	< 1.6 $\times 10^{-4}$	CL=95%
$\Gamma_{47}$ $\pi^- K^0 \overline{K}^0 \nu_\tau$	( 1.55 $\pm$ 0.24 ) $\times 10^{-3}$	
$\Gamma_{48}$ $\pi^- K_S^0 K_S^0 \nu_\tau$	[a] ( 2.33 $\pm$ 0.07 ) $\times 10^{-4}$	
$\Gamma_{49}$ $\pi^- K_S^0 K_L^0 \nu_\tau$	[a] ( 1.08 $\pm$ 0.24 ) $\times 10^{-3}$	
$\Gamma_{50}$ $\pi^- K_L^0 K_L^0 \nu_\tau$	( 2.33 $\pm$ 0.07 ) $\times 10^{-4}$	
$\Gamma_{51}$ $\pi^- K^0 \overline{K}^0 \pi^0 \nu_\tau$	( 3.6 $\pm$ 1.2 ) $\times 10^{-4}$	
$\Gamma_{52}$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	[a] ( 1.82 $\pm$ 0.21 ) $\times 10^{-5}$	
$\Gamma_{53}$ $K^{*-} K^0 \pi^0 \nu_\tau \rightarrow$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	( 1.08 $\pm$ 0.21 ) $\times 10^{-5}$	

$\Gamma_{54}$	$f_1(1285) \pi^- \nu_\tau \rightarrow$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	( 6.8 $\pm$ 1.5 ) $\times 10^{-6}$	
$\Gamma_{55}$	$f_1(1420) \pi^- \nu_\tau \rightarrow$ $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	( 2.4 $\pm$ 0.8 ) $\times 10^{-6}$	
$\Gamma_{56}$	$\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau$	[a] ( 3.2 $\pm$ 1.2 ) $\times 10^{-4}$	
$\Gamma_{57}$	$\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau$	( 1.82 $\pm$ 0.21 ) $\times 10^{-5}$	
$\Gamma_{58}$	$K^- K_S^0 K_S^0 \nu_\tau$	< 6.3 $\times 10^{-7}$	CL=90%
$\Gamma_{59}$	$K^- K_S^0 K_S^0 \pi^0 \nu_\tau$	< 4.0 $\times 10^{-7}$	CL=90%
$\Gamma_{60}$	$K^0 h^+ h^- h^- \geq 0$ neutrals $\nu_\tau$	< 1.7 $\times 10^{-3}$	CL=95%
$\Gamma_{61}$	$K^0 h^+ h^- h^- \nu_\tau$	[a] ( 2.5 $\pm$ 2.0 ) $\times 10^{-4}$	

## Modes with three charged particles

$\Gamma_{62}$	$h^- h^- h^+ \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	(15.21 $\pm$ 0.06) %	
$\Gamma_{63}$	$h^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$ ) (“3-prong”)	(14.55 $\pm$ 0.06) %	
$\Gamma_{64}$	$h^- h^- h^+ \nu_\tau$	( 9.80 $\pm$ 0.05 ) %	
$\Gamma_{65}$	$h^- h^- h^+ \nu_\tau$ (ex. $K^0$ )	( 9.46 $\pm$ 0.05 ) %	
$\Gamma_{66}$	$h^- h^- h^+ \nu_\tau$ (ex. $K^0, \omega$ )	( 9.43 $\pm$ 0.05 ) %	
$\Gamma_{67}$	$\pi^- \pi^+ \pi^- \nu_\tau$	( 9.31 $\pm$ 0.05 ) %	
$\Gamma_{68}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	( 9.02 $\pm$ 0.05 ) %	
$\Gamma_{69}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ), non-axial vector	< 2.4 %	CL=95%
$\Gamma_{70}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[a] ( 8.99 $\pm$ 0.05 ) %	
$\Gamma_{71}$	$h^- h^- h^+ \geq 1$ neutrals $\nu_\tau$	( 5.29 $\pm$ 0.05 ) %	
$\Gamma_{72}$	$h^- h^- h^+ \geq 1 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 5.09 $\pm$ 0.05 ) %	
$\Gamma_{73}$	$h^- h^- h^+ \pi^0 \nu_\tau$	( 4.76 $\pm$ 0.05 ) %	
$\Gamma_{74}$	$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	( 4.57 $\pm$ 0.05 ) %	
$\Gamma_{75}$	$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	( 2.79 $\pm$ 0.07 ) %	
$\Gamma_{76}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$	( 4.62 $\pm$ 0.05 ) %	
$\Gamma_{77}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )	( 4.49 $\pm$ 0.05 ) %	
$\Gamma_{78}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	[a] ( 2.74 $\pm$ 0.07 ) %	
$\Gamma_{79}$	$h^- \rho \pi^0 \nu_\tau$		
$\Gamma_{80}$	$h^- \rho^+ h^- \nu_\tau$		
$\Gamma_{81}$	$h^- \rho^- h^+ \nu_\tau$		
$\Gamma_{82}$	$h^- h^- h^+ \geq 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 5.17 $\pm$ 0.31 ) $\times 10^{-3}$	
$\Gamma_{83}$	$h^- h^- h^+ 2 \pi^0 \nu_\tau$	( 5.05 $\pm$ 0.31 ) $\times 10^{-3}$	
$\Gamma_{84}$	$h^- h^- h^+ 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 4.95 $\pm$ 0.31 ) $\times 10^{-3}$	
$\Gamma_{85}$	$h^- h^- h^+ 2 \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[a] (10 $\pm$ 4 ) $\times 10^{-4}$	
$\Gamma_{86}$	$h^- h^- h^+ 3 \pi^0 \nu_\tau$	( 2.12 $\pm$ 0.30 ) $\times 10^{-4}$	
$\Gamma_{87}$	$2 \pi^- \pi^+ 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.94 $\pm$ 0.30 ) $\times 10^{-4}$	
$\Gamma_{88}$	$2 \pi^- \pi^+ 3 \pi^0 \nu_\tau$ (ex. $K^0, \eta$ , $f_1(1285)$ )	( 1.7 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{89}$	$2 \pi^- \pi^+ 3 \pi^0 \nu_\tau$ (ex. $K^0, \eta$ , $\omega, f_1(1285)$ )	[a] ( 1.4 $\pm$ 2.7 ) $\times 10^{-5}$	
$\Gamma_{90}$	$K^- h^+ h^- \geq 0$ neutrals $\nu_\tau$	( 6.29 $\pm$ 0.14 ) $\times 10^{-3}$	
$\Gamma_{91}$	$K^- h^+ \pi^- \nu_\tau$ (ex. $K^0$ )	( 4.37 $\pm$ 0.07 ) $\times 10^{-3}$	
$\Gamma_{92}$	$K^- h^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )	( 8.6 $\pm$ 1.2 ) $\times 10^{-4}$	
$\Gamma_{93}$	$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$	( 4.77 $\pm$ 0.14 ) $\times 10^{-3}$	
$\Gamma_{94}$	$K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 3.73 $\pm$ 0.13 ) $\times 10^{-3}$	
$\Gamma_{95}$	$K^- \pi^+ \pi^- \nu_\tau$	( 3.45 $\pm$ 0.07 ) $\times 10^{-3}$	
$\Gamma_{96}$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	( 2.93 $\pm$ 0.07 ) $\times 10^{-3}$	
$\Gamma_{97}$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[a] ( 2.93 $\pm$ 0.07 ) $\times 10^{-3}$	
$\Gamma_{98}$	$K^- \rho^0 \nu_\tau \rightarrow$ $K^- \pi^+ \pi^- \nu_\tau$	( 1.4 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{99}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$	( 1.31 $\pm$ 0.12 ) $\times 10^{-3}$	
$\Gamma_{100}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )	( 7.9 $\pm$ 1.2 ) $\times 10^{-4}$	
$\Gamma_{101}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	( 7.6 $\pm$ 1.2 ) $\times 10^{-4}$	
$\Gamma_{102}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	( 3.7 $\pm$ 0.9 ) $\times 10^{-4}$	
$\Gamma_{103}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[a] ( 3.9 $\pm$ 1.4 ) $\times 10^{-4}$	
$\Gamma_{104}$	$K^- \pi^+ K^- \geq 0$ neut. $\nu_\tau$	< 9 $\times 10^{-4}$	CL=95%
$\Gamma_{105}$	$K^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$	( 1.496 $\pm$ 0.033 ) $\times 10^{-3}$	
$\Gamma_{106}$	$K^- K^+ \pi^- \nu_\tau$	[a] ( 1.435 $\pm$ 0.027 ) $\times 10^{-3}$	
$\Gamma_{107}$	$K^- K^+ \pi^- \pi^0 \nu_\tau$	[a] ( 6.1 $\pm$ 1.8 ) $\times 10^{-5}$	
$\Gamma_{108}$	$K^- K^+ K^- \nu_\tau$	( 2.2 $\pm$ 0.8 ) $\times 10^{-5}$	S=5.4
$\Gamma_{109}$	$K^- K^+ K^- \nu_\tau$ (ex. $\phi$ )	< 2.5 $\times 10^{-6}$	CL=90%
$\Gamma_{110}$	$K^- K^+ K^- \pi^0 \nu_\tau$	< 4.8 $\times 10^{-6}$	CL=90%
$\Gamma_{111}$	$\pi^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$	< 2.5 $\times 10^{-3}$	CL=95%
$\Gamma_{112}$	$e^- e^- e^+ \overline{\nu}_e \nu_\tau$	( 2.8 $\pm$ 1.5 ) $\times 10^{-5}$	
$\Gamma_{113}$	$\mu^- e^- e^+ \overline{\nu}_\mu \nu_\tau$	< 3.6 $\times 10^{-5}$	CL=90%

See key on page 601

## Lepton Particle Listings

T

## Modes with five charged particles

$\Gamma_{114}$	$3h^-2h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^- \pi^+$ ) ("5-prong")	( 9.9 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{115}$	$3h^-2h^+ \nu_\tau$ (ex. $K^0$ )	( 8.22 $\pm$ 0.32 ) $\times 10^{-4}$	
$\Gamma_{116}$	$3\pi^-2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ )	( 8.21 $\pm$ 0.31 ) $\times 10^{-4}$	
$\Gamma_{117}$	$3\pi^-2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ , $f_1(1285)$ )	[a] ( 7.69 $\pm$ 0.30 ) $\times 10^{-4}$	
$\Gamma_{118}$	$K^-2\pi^-2\pi^+ \nu_\tau$ (ex. $K^0$ )	[a] ( 6 $\pm$ 12 ) $\times 10^{-7}$	
$\Gamma_{119}$	$K^+3\pi^- \pi^+ \nu_\tau$	< 5.0 $\times 10^{-6}$	CL=90%
$\Gamma_{120}$	$K^+K^-2\pi^- \pi^+ \nu_\tau$	< 4.5 $\times 10^{-7}$	CL=90%
$\Gamma_{121}$	$3h^-2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.64 $\pm$ 0.11 ) $\times 10^{-4}$	
$\Gamma_{122}$	$3\pi^-2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.62 $\pm$ 0.11 ) $\times 10^{-4}$	
$\Gamma_{123}$	$3\pi^-2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta$ , $f_1(1285)$ )	( 1.11 $\pm$ 0.10 ) $\times 10^{-4}$	
$\Gamma_{124}$	$3\pi^-2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega$ , $f_1(1285)$ )	[a] ( 3.8 $\pm$ 0.9 ) $\times 10^{-5}$	
$\Gamma_{125}$	$K^-2\pi^-2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 1.1 $\pm$ 0.6 ) $\times 10^{-6}$	
$\Gamma_{126}$	$K^+3\pi^- \pi^+ \pi^0 \nu_\tau$	< 8 $\times 10^{-7}$	CL=90%
$\Gamma_{127}$	$3h^-2h^+2\pi^0 \nu_\tau$	< 3.4 $\times 10^{-6}$	CL=90%

## Miscellaneous other allowed modes

$\Gamma_{128}$	$(5\pi)^- \nu_\tau$	( 7.8 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{129}$	$4h^-3h^+ \geq 0$ neutrals $\nu_\tau$ ("7-prong")	< 3.0 $\times 10^{-7}$	CL=90%
$\Gamma_{130}$	$4h^-3h^+ \nu_\tau$	< 4.3 $\times 10^{-7}$	CL=90%
$\Gamma_{131}$	$4h^-3h^+ \pi^0 \nu_\tau$	< 2.5 $\times 10^{-7}$	CL=90%
$\Gamma_{132}$	$X^-(S=-1) \nu_\tau$	( 2.92 $\pm$ 0.04 ) %	
$\Gamma_{133}$	$K^*(892)^- \geq 0$ neutrals $\geq$ $0 K_L^0 \nu_\tau$	( 1.42 $\pm$ 0.18 ) %	S=1.4
$\Gamma_{134}$	$K^*(892)^- \nu_\tau$	( 1.20 $\pm$ 0.07 ) %	S=1.8
$\Gamma_{135}$	$K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau$	( 7.83 $\pm$ 0.26 ) $\times 10^{-3}$	
$\Gamma_{136}$	$K^*(892)^0 K^- \geq 0$ neutrals $\nu_\tau$	( 3.2 $\pm$ 1.4 ) $\times 10^{-3}$	
$\Gamma_{137}$	$K^*(892)^0 K^- \nu_\tau$	( 2.1 $\pm$ 0.4 ) $\times 10^{-3}$	
$\Gamma_{138}$	$\bar{K}^*(892)^0 \pi^- \geq 0$ neutrals $\nu_\tau$	( 3.8 $\pm$ 1.7 ) $\times 10^{-3}$	
$\Gamma_{139}$	$\bar{K}^*(892)^0 \pi^- \nu_\tau$	( 2.2 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{140}$	$(\bar{K}^*(892) \pi)^- \nu_\tau \rightarrow$ $\pi^- \bar{K}^0 \pi^0 \nu_\tau$	( 1.0 $\pm$ 0.4 ) $\times 10^{-3}$	
$\Gamma_{141}$	$K_1(1270)^- \nu_\tau$	( 4.7 $\pm$ 1.1 ) $\times 10^{-3}$	
$\Gamma_{142}$	$K_1(1400)^- \nu_\tau$	( 1.7 $\pm$ 2.6 ) $\times 10^{-3}$	S=1.7
$\Gamma_{143}$	$K^*(1410)^- \nu_\tau$	( 1.5 $\pm$ 1.4 ) $\times 10^{-3}$	
$\Gamma_{144}$	$K_0^*(1430)^- \nu_\tau$	< 5 $\times 10^{-4}$	CL=95%
$\Gamma_{145}$	$K_2^*(1430)^- \nu_\tau$	< 3 $\times 10^{-3}$	CL=95%
$\Gamma_{146}$	$a_0(980)^- \geq 0$ neutrals $\nu_\tau$		
$\Gamma_{147}$	$\eta \pi^- \nu_\tau$	< 9.9 $\times 10^{-5}$	CL=95%
$\Gamma_{148}$	$\eta \pi^- \pi^0 \nu_\tau$	[a] ( 1.39 $\pm$ 0.07 ) $\times 10^{-3}$	
$\Gamma_{149}$	$\eta \pi^- \pi^0 \pi^0 \nu_\tau$	[a] ( 1.9 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{150}$	$\eta K^- \nu_\tau$	[a] ( 1.55 $\pm$ 0.08 ) $\times 10^{-4}$	
$\Gamma_{151}$	$\eta K^*(892)^- \nu_\tau$	( 1.38 $\pm$ 0.15 ) $\times 10^{-4}$	
$\Gamma_{152}$	$\eta K^- \pi^0 \nu_\tau$	[a] ( 4.8 $\pm$ 1.2 ) $\times 10^{-5}$	
$\Gamma_{153}$	$\eta K^- \pi^0$ (non- $K^*(892)$ ) $\nu_\tau$	< 3.5 $\times 10^{-5}$	CL=90%
$\Gamma_{154}$	$\eta \bar{K}^0 \pi^- \nu_\tau$	[a] ( 9.4 $\pm$ 1.5 ) $\times 10^{-5}$	
$\Gamma_{155}$	$\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau$	< 5.0 $\times 10^{-5}$	CL=90%
$\Gamma_{156}$	$\eta K^- K^0 \nu_\tau$	< 9.0 $\times 10^{-6}$	CL=90%
$\Gamma_{157}$	$\eta \pi^+ \pi^- \pi^- \geq 0$ neutrals $\nu_\tau$	< 3 $\times 10^{-3}$	CL=90%
$\Gamma_{158}$	$\eta \pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	[a] ( 2.19 $\pm$ 0.13 ) $\times 10^{-4}$	
$\Gamma_{159}$	$\eta \pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, f_1(1285)$ )	( 9.9 $\pm$ 1.6 ) $\times 10^{-5}$	
$\Gamma_{160}$	$\eta a_1(1260)^- \nu_\tau \rightarrow \eta \pi^- \rho^0 \nu_\tau$	< 3.9 $\times 10^{-4}$	CL=90%
$\Gamma_{161}$	$\eta \eta \pi^- \nu_\tau$	< 7.4 $\times 10^{-6}$	CL=90%
$\Gamma_{162}$	$\eta \eta \pi^- \pi^0 \nu_\tau$	< 2.0 $\times 10^{-4}$	CL=95%
$\Gamma_{163}$	$\eta \eta K^- \nu_\tau$	< 3.0 $\times 10^{-6}$	CL=90%
$\Gamma_{164}$	$\eta'(958) \pi^- \nu_\tau$	< 4.0 $\times 10^{-6}$	CL=90%
$\Gamma_{165}$	$\eta'(958) \pi^- \pi^0 \nu_\tau$	< 1.2 $\times 10^{-5}$	CL=90%
$\Gamma_{166}$	$\eta'(958) K^- \nu_\tau$	< 2.4 $\times 10^{-6}$	CL=90%
$\Gamma_{167}$	$\phi \pi^- \nu_\tau$	( 3.4 $\pm$ 0.6 ) $\times 10^{-5}$	
$\Gamma_{168}$	$\phi K^- \nu_\tau$	[a] ( 4.4 $\pm$ 1.6 ) $\times 10^{-5}$	
$\Gamma_{169}$	$f_1(1285) \pi^- \nu_\tau$	( 3.9 $\pm$ 0.5 ) $\times 10^{-4}$	S=1.9
$\Gamma_{170}$	$f_1(1285) \pi^- \nu_\tau \rightarrow$ $\eta \pi^- \pi^+ \pi^- \nu_\tau$	( 1.18 $\pm$ 0.07 ) $\times 10^{-4}$	S=1.3
$\Gamma_{171}$	$f_1(1285) \pi^- \nu_\tau \rightarrow 3\pi^-2\pi^+ \nu_\tau$	[a] ( 5.2 $\pm$ 0.4 ) $\times 10^{-5}$	
$\Gamma_{172}$	$\pi(1300)^- \nu_\tau \rightarrow (\rho \pi)^- \nu_\tau \rightarrow$ ( $3\pi$ ) $^- \nu_\tau$	< 1.0 $\times 10^{-4}$	CL=90%
$\Gamma_{173}$	$\pi(1300)^- \nu_\tau \rightarrow$ ( $\pi \pi$ ) $S$ -wave $\pi^- \nu_\tau \rightarrow$ ( $3\pi$ ) $^- \nu_\tau$	< 1.9 $\times 10^{-4}$	CL=90%

$\Gamma_{174}$	$h^- \omega \geq 0$ neutrals $\nu_\tau$	( 2.40 $\pm$ 0.08 ) %	
$\Gamma_{175}$	$h^- \omega \nu_\tau$	( 1.99 $\pm$ 0.06 ) %	
$\Gamma_{176}$	$\pi^- \omega \nu_\tau$	[a] ( 1.95 $\pm$ 0.06 ) %	
$\Gamma_{177}$	$K^- \omega \nu_\tau$	[a] ( 4.1 $\pm$ 0.9 ) $\times 10^{-4}$	
$\Gamma_{178}$	$h^- \omega \pi^0 \nu_\tau$	[a] ( 4.1 $\pm$ 0.4 ) $\times 10^{-3}$	
$\Gamma_{179}$	$h^- \omega 2\pi^0 \nu_\tau$	( 1.4 $\pm$ 0.5 ) $\times 10^{-4}$	
$\Gamma_{180}$	$\pi^- \omega 2\pi^0 \nu_\tau$	[a] ( 7.1 $\pm$ 1.6 ) $\times 10^{-5}$	
$\Gamma_{181}$	$h^- 2\omega \nu_\tau$	< 5.4 $\times 10^{-7}$	CL=90%
$\Gamma_{182}$	$2h^- h^+ \omega \nu_\tau$	( 1.20 $\pm$ 0.22 ) $\times 10^{-4}$	
$\Gamma_{183}$	$2\pi^- \pi^+ \omega \nu_\tau$ (ex. $K^0$ )	[a] ( 8.4 $\pm$ 0.6 ) $\times 10^{-5}$	

Lepton Family number (LF), Lepton number (L),  
or Baryon number (B) violating modes

L means lepton number violation (e.g.  $\tau^- \rightarrow e^+ \pi^- \pi^-$ ). Following common usage, LF means lepton family violation and not lepton number violation (e.g.  $\tau^- \rightarrow e^- \pi^+ \pi^-$ ). B means baryon number violation.

$\Gamma_{184}$	$e^- \gamma$	LF	< 3.3	$\times 10^{-8}$	CL=90%
$\Gamma_{185}$	$\mu^- \gamma$	LF	< 4.4	$\times 10^{-8}$	CL=90%
$\Gamma_{186}$	$e^- \pi^0$	LF	< 8.0	$\times 10^{-8}$	CL=90%
$\Gamma_{187}$	$\mu^- \pi^0$	LF	< 1.1	$\times 10^{-7}$	CL=90%
$\Gamma_{188}$	$e^- K_S^0$	LF	< 2.6	$\times 10^{-8}$	CL=90%
$\Gamma_{189}$	$\mu^- K_S^0$	LF	< 2.3	$\times 10^{-8}$	CL=90%
$\Gamma_{190}$	$e^- \eta$	LF	< 9.2	$\times 10^{-8}$	CL=90%
$\Gamma_{191}$	$\mu^- \eta$	LF	< 6.5	$\times 10^{-8}$	CL=90%
$\Gamma_{192}$	$e^- \rho^0$	LF	< 1.8	$\times 10^{-8}$	CL=90%
$\Gamma_{193}$	$\mu^- \rho^0$	LF	< 1.2	$\times 10^{-8}$	CL=90%
$\Gamma_{194}$	$e^- \omega$	LF	< 4.8	$\times 10^{-8}$	CL=90%
$\Gamma_{195}$	$\mu^- \omega$	LF	< 4.7	$\times 10^{-8}$	CL=90%
$\Gamma_{196}$	$e^- K^*(892)^0$	LF	< 3.2	$\times 10^{-8}$	CL=90%
$\Gamma_{197}$	$\mu^- K^*(892)^0$	LF	< 5.9	$\times 10^{-8}$	CL=90%
$\Gamma_{198}$	$e^- \bar{K}^*(892)^0$	LF	< 3.4	$\times 10^{-8}$	CL=90%
$\Gamma_{199}$	$\mu^- \bar{K}^*(892)^0$	LF	< 7.0	$\times 10^{-8}$	CL=90%
$\Gamma_{200}$	$e^- \eta'(958)$	LF	< 1.6	$\times 10^{-7}$	CL=90%
$\Gamma_{201}$	$\mu^- \eta'(958)$	LF	< 1.3	$\times 10^{-7}$	CL=90%
$\Gamma_{202}$	$e^- f_0(980) \rightarrow e^- \pi^+ \pi^-$	LF	< 3.2	$\times 10^{-8}$	CL=90%
$\Gamma_{203}$	$\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-$	LF	< 3.4	$\times 10^{-8}$	CL=90%
$\Gamma_{204}$	$e^- \phi$	LF	< 3.1	$\times 10^{-8}$	CL=90%
$\Gamma_{205}$	$\mu^- \phi$	LF	< 8.4	$\times 10^{-8}$	CL=90%
$\Gamma_{206}$	$e^- e^+ e^-$	LF	< 2.7	$\times 10^{-8}$	CL=90%
$\Gamma_{207}$	$e^- \mu^+ \mu^-$	LF	< 2.7	$\times 10^{-8}$	CL=90%
$\Gamma_{208}$	$e^+ \mu^- \mu^-$	LF	< 1.7	$\times 10^{-8}$	CL=90%
$\Gamma_{209}$	$\mu^- e^+ e^-$	LF	< 1.8	$\times 10^{-8}$	CL=90%
$\Gamma_{210}$	$\mu^+ e^- e^-$	LF	< 1.5	$\times 10^{-8}$	CL=90%
$\Gamma_{211}$	$\mu^- \mu^+ \mu^-$	LF	< 2.1	$\times 10^{-8}$	CL=90%
$\Gamma_{212}$	$e^- \pi^+ \pi^-$	LF	< 2.3	$\times 10^{-8}$	CL=90%
$\Gamma_{213}$	$e^+ \pi^- \pi^-$	L	< 2.0	$\times 10^{-8}$	CL=90%
$\Gamma_{214}$	$\mu^- \pi^+ \pi^-$	LF	< 2.1	$\times 10^{-8}$	CL=90%
$\Gamma_{215}$	$\mu^+ \pi^- \pi^-$	L	< 3.9	$\times 10^{-8}$	CL=90%
$\Gamma_{216}$	$e^- \pi^+ K^-$	LF	< 3.7	$\times 10^{-8}$	CL=90%
$\Gamma_{217}$	$e^- \pi^- K^+$	LF	< 3.1	$\times 10^{-8}$	CL=90%
$\Gamma_{218}$	$e^+ \pi^- K^-$	L	< 3.2	$\times 10^{-8}$	CL=90%
$\Gamma_{219}$	$e^- K_S^0 K_S^0$	LF	< 7.1	$\times 10^{-8}$	CL=90%
$\Gamma_{220}$	$e^- K^+ K^-$	LF	< 3.4	$\times 10^{-8}$	CL=90%
$\Gamma_{221}$	$e^+ K^- K^-$	L	< 3.3	$\times 10^{-8}$	CL=90%
$\Gamma_{222}$	$\mu^- \pi^+ K^-$	LF	< 8.6	$\times 10^{-8}$	CL=90%
$\Gamma_{223}$	$\mu^- \pi^- K^+$	LF	< 4.5	$\times 10^{-8}$	CL=90%
$\Gamma_{224}$	$\mu^+ \pi^- K^-$	L	< 4.8	$\times 10^{-8}$	CL=90%
$\Gamma_{225}$	$\mu^- K_S^0 K_S^0$	LF	< 8.0	$\times 10^{-8}$	CL=90%
$\Gamma_{226}$	$\mu^- K^+ K^-$	LF	< 4.4	$\times 10^{-8}$	CL=90%
$\Gamma_{227}$	$\mu^+ K^- K^-$	L	< 4.7	$\times 10^{-8}$	CL=90%
$\Gamma_{228}$	$e^- \pi^0 \pi^0$	LF	< 6.5	$\times 10^{-6}$	CL=90%
$\Gamma_{229}$	$\mu^- \pi^0 \pi^0$	LF	< 1.4	$\times 10^{-5}$	CL=90%
$\Gamma_{230}$	$e^- \eta \eta$	LF	< 3.5	$\times 10^{-5}$	CL=90%
$\Gamma_{231}$	$\mu^- \eta \eta$	LF	< 6.0	$\times 10^{-5}$	CL=90%
$\Gamma_{232}$	$e^- \pi^0 \eta$	LF	< 2.4	$\times 10^{-5}$	CL=90%
$\Gamma_{233}$	$\mu^- \pi^0 \eta$	LF	< 2.2	$\times 10^{-5}$	CL=90%
$\Gamma_{234}$	$\rho \mu^- \mu^-$	L,B	< 4.4	$\times 10^{-7}$	CL=90%
$\Gamma_{235}$	$\bar{\rho} \mu^+ \mu^-$	L,B	< 3.3	$\times 10^{-7}$	CL=90%
$\Gamma_{236}$	$\bar{\rho} \gamma$	L,B	< 3.5	$\times 10^{-6}$	CL=90%
$\Gamma_{237}$	$\bar{\rho} \pi^0$	L,B	< 1.5	$\times 10^{-5}$	CL=90%
$\Gamma_{238}$	$\bar{\rho} 2\pi^0$	L,B	< 3.3	$\times 10^{-5}$	CL=90%
$\Gamma_{239}$	$\bar{\rho} \eta$	L,B	< 8.9	$\times 10^{-6}$	CL=90%
$\Gamma_{240}$	$\bar{\rho} \pi^0 \eta$	L,B	< 2.7	$\times 10^{-5}$	CL=90%
$\Gamma_{241}$	$\Lambda \pi^-$	L,B	< 7.2	$\times 10^{-8}$	CL=90%

$\Gamma_{242}$	$\bar{L}\pi^-$	$L, B$	$< 1.4$	$\times 10^{-7}$	CL=90%
$\Gamma_{243}$	$e^- \text{ light boson}$	$LF$	$< 2.7$	$\times 10^{-3}$	CL=95%
$\Gamma_{244}$	$\mu^- \text{ light boson}$	$LF$	$< 5$	$\times 10^{-3}$	CL=95%

[b] See the Particle Listings below for the energy limits used in this measurement.

An overall fit to 85 branching ratios uses 169 measurements and one constraint to determine 46 parameters. The overall fit has a  $\chi^2 = 134.9$  for 124 degrees of freedom.

[illegible]

$x_{36}$	0									
$x_{38}$	0	-22								
$x_{41}$	0	-13	4							
$x_{43}$	0	2	-21	-20						
$x_{45}$	0	-3	0	-6	0					
$x_{48}$	0	-1	1	-4	1	0				
$x_{49}$	0	-5	0	-4	-1	-10	0			
$x_{52}$	0	0	7	0	5	0	-7	0		
$x_{56}$	0	-2	0	-2	-1	-4	0	-8	0	
$x_{61}$	0	-2	0	-2	0	-4	0	-4	0	-2
$x_{70}$	-5	-2	0	-1	0	-4	1	-4	0	-2
$x_{78}$	3	1	0	1	0	2	0	2	0	1
$x_{85}$	2	0	0	0	0	0	0	0	0	0
$x_{89}$	0	0	0	0	0	0	0	0	-1	0
$x_{97}$	-1	-1	0	-1	0	-2	0	-2	0	-1
$x_{103}$	-1	-1	0	-1	0	-1	0	-1	0	-1
$x_{106}$	-1	-1	0	0	0	-1	0	-1	0	0
$x_{107}$	0	0	0	0	0	0	0	0	0	0
$x_{117}$	-1	0	0	0	0	-1	0	-1	0	0
$x_{118}$	0	0	0	0	0	0	0	0	0	0
$x_{124}$	0	0	0	0	0	0	0	0	0	0
$x_{125}$	0	0	0	0	0	0	0	0	0	0
$x_{148}$	-2	0	0	0	0	-1	1	-1	0	0
$x_{149}$	0	0	0	0	0	-1	0	-1	0	0
$x_{150}$	0	0	0	0	0	0	1	0	0	0
$x_{152}$	0	0	0	0	0	0	0	0	0	0
$x_{154}$	0	0	0	0	0	0	0	-1	0	0
$x_{158}$	-1	0	0	0	0	-1	0	-1	0	0
$x_{168}$	0	0	0	0	0	0	0	0	0	0
$x_{171}$	0	0	0	0	0	0	0	0	0	0
$x_{176}$	1	-1	0	0	0	-1	0	-1	0	0
$x_{177}$	0	0	0	0	0	0	0	0	0	0
$x_{178}$	2	-1	0	0	0	-1	0	-1	0	0
$x_{180}$	0	0	0	0	0	0	0	0	0	0
$x_{183}$	-1	0	0	0	0	0	0	0	0	0
	$x_{30}$	$x_{36}$	$x_{38}$	$x_{41}$	$x_{43}$	$x_{45}$	$x_{48}$	$x_{49}$	$x_{52}$	$x_{56}$
$x_{70}$	-4									
$x_{78}$	2	-19								
$x_{85}$	0	-1	-8							
$x_{89}$	0	-1	-1	0						
$x_{97}$	-2	19	-6	0	0					
$x_{103}$	-1	-4	-14	-1	0	-1				
$x_{106}$	-1	15	-4	0	0	0	-1			
$x_{107}$	0	-1	-1	0	0	0	-3	0		
$x_{117}$	-1	0	0	0	-4	0	0	0	0	
$x_{118}$	0	0	0	0	0	0	0	0	0	-1
$x_{124}$	0	0	0	0	0	0	0	0	0	3
$x_{125}$	0	0	0	0	0	0	0	0	0	-1
$x_{148}$	-1	0	0	-5	0	0	0	0	0	0
$x_{149}$	-1	-1	0	0	-11	0	0	0	0	10
$x_{150}$	0	2	0	0	0	0	-1	1	0	0
$x_{152}$	0	0	0	-1	0	0	0	0	0	0
$x_{154}$	0	0	0	0	-2	0	0	0	0	0
$x_{158}$	-1	-1	0	0	-8	0	0	0	0	47
$x_{168}$	0	-1	0	0	0	1	0	1	0	0</



$x_{124}$	0									
$x_{125}$	0	-1								
$x_{148}$	0	0	0							
$x_{149}$	0	2	0	0						
$x_{150}$	0	0	0	4						
$x_{152}$	0	0	0	1	0	1				
$x_{154}$	0	0	0	2	-1	1	0			
$x_{158}$	-1	3	-1	0	25	0	0	0		
$x_{168}$	0	0	0	0	0	0	0	0	0	
$x_{171}$	-1	1	0	0	4	0	0	0	20	0
$x_{176}$	0	0	0	0	0	0	0	0	0	0
$x_{177}$	0	0	0	0	0	0	0	0	0	0
$x_{178}$	0	0	0	0	0	0	0	0	0	0
$x_{180}$	0	2	0	0	10	0	0	-1	20	0
$x_{183}$	-1	-2	-1	0	17	0	0	0	39	0
	$x_{118}$	$x_{124}$	$x_{125}$	$x_{148}$	$x_{149}$	$x_{150}$	$x_{152}$	$x_{154}$	$x_{158}$	$x_{168}$

$x_{176}$	0				
$x_{177}$	0	-14			
$x_{178}$	0	-4	0		
$x_{180}$	3	0	0	0	
$x_{183}$	17	0	0	0	14
	$x_{171}$	$x_{176}$	$x_{177}$	$x_{178}$	$x_{180}$

### $\tau$ BRANCHING FRACTIONS

Revised April 2016 by S.Banerjee (University of Louisville), K.Hayes (Hillsdale College), A.Lusiani (Scuola Normale Superiore and INFN, sezione di Pisa)

In order to make optimal use of the experimental data to determine the  $\tau$  branching fractions, their uncertainties, and their correlations, we perform a global minimum  $\chi^2$  fit using the measured values, their uncertainties, their statistical correlations, their dependencies on external parameters and common systematics, and the relations that hold between the branching fractions, including a unitarity constraint on the sum of all the exclusive  $\tau$  decay branching fractions. Starting with this edition, we use a new fit procedure, which has been elaborated by the Tau Physics Group within the Heavy Flavour Averaging Group (HFAG) [1].

In the following, we use “branching fraction” to refer to the partial decay fraction of a particle like the  $\tau$  into a specific decay mode, and “branching ratio” to refer to quantities derived from the branching fractions [2], like for instance a ratio of two branching fractions, or a ratio of two linear combinations of branching fractions.

#### *The constrained fit to $\tau$ branching fractions.*

The  $\tau$  Listings contains 242  $\tau$  decay modes, out of which 61 are Lepton Family number, Lepton number, or Baryon number violating modes. The fit computes the branching fractions of 112 decay modes. Although no new  $\tau$  branching fraction and ratio measurements have been released since the 2015 edition, the fit in this edition includes more experimental measurements (169, up from 143 in 2015) and determines in the fit several additional  $\tau$  branching fractions and ratios, relying on a larger and updated set of constraints that relate the branching fractions and ratios between themselves. The measurements are treated as follows [1].

Many published measurements depend on external parameters such as the  $\tau$  pair production cross-section in  $e^+e^-$

annihilations at the  $\Upsilon(4S)$  peak. We compute the size and sign of these dependencies and update the measurements and their uncertainties to the current values of the external parameters. Accordingly, the measurements and their uncertainties are updated to account for updated values of external parameters. The dependencies on common systematic effects are also determined in size and sign, and all the common systematic dependencies of different measurements are used together with the published statistical and systematic uncertainties and correlations in order to compute a single all-inclusive variance and covariance matrix of the experimental measurements. All the measurements, their uncertainties, and their correlations were taken from the respective published papers. Their values and the constraints used in the fit are reported in the  $\tau$  Listings section that follows this review. If only a few measurements are correlated, the correlation coefficients are listed in the footnote for each measurement (see for example  $\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau)$  (“1-prong”))/ $\Gamma_{\text{total}}$ ). If a large number of measurements are correlated, then the full correlation matrix is listed in the footnote to the measurement that first appears in the  $\tau$  Listings. Footnotes to the other measurements refer to the first measurement. For example, the large correlation matrices for the branching fraction or ratio measurements contained in Refs. [3,4] are listed in Footnotes to the  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$  measurements respectively. The constraints between the  $\tau$  branching fractions and ratios include coefficients that correspond to physical quantities, like for instance the branching fractions of the  $\eta$  and  $\omega$  mesons. All quantities are taken from the 2015 edition of the Review of Particle Physics. Their uncertainties are neglected in the fit.

Compared to the 2015 edition, the fit now includes several additional modes, mainly related to the most recent BaBar papers on high multiplicity modes [5] and  $K_S^0 K_S^0$  modes [6] and the Belle paper on neutral kaon modes [7]:

$$B(\tau \rightarrow \pi^- \pi^0 K_S^0 K_S^0 \nu_\tau)$$

$$B(\tau \rightarrow K^- K^- K^+ \nu_\tau)$$

$$B(\tau \rightarrow K^- \pi^0 \eta \nu_\tau)$$

$$B(\tau \rightarrow \pi^- \bar{K}^0 \eta \nu_\tau) ;$$

Also, the following components of  $\tau$ -decay modes are now included [5,8,9]:

$$B(\tau \rightarrow \pi^- 2\pi^0 \eta \nu_\tau \text{ } (\eta \rightarrow \pi^+ \pi^- \pi^0) \text{ (ex. } K^0))$$

$$B(\tau \rightarrow 2\pi^- \pi^+ \eta \nu_\tau \text{ } (\eta \rightarrow \pi^+ \pi^- \pi^0) \text{ (ex. } K^0))$$

$$B(\tau \rightarrow 2\pi^- \pi^+ \eta \nu_\tau \text{ } (\eta \rightarrow \gamma \gamma) \text{ (ex. } K^0))$$

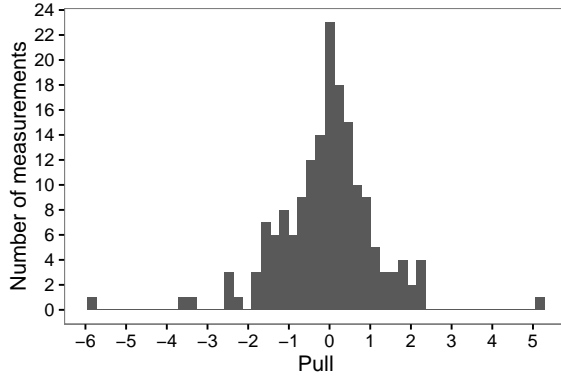
$$B(\tau \rightarrow \pi^- 2\pi^0 \omega \nu_\tau \text{ (ex. } K^0))$$

$$B(\tau \rightarrow 2\pi^- \pi^+ \omega \nu_\tau \text{ (ex. } K^0))$$

$$B(\tau \rightarrow \pi^- f_1 \nu_\tau \text{ } (f_1 \rightarrow 2\pi^- 2\pi^+)) .$$

$$B(\tau \rightarrow K^- \phi \nu_\tau) .$$

We obtain the branching fraction of  $\tau \rightarrow a_1^- (\rightarrow \pi^- \gamma) \nu_\tau$  using the ALEPH estimate for  $B(a_1^- \rightarrow \pi^- \gamma)$  [3], which uses the measurement of  $\Gamma(a_1^- \rightarrow \pi^- \gamma)$  [10]. In the fit, we assume that  $B(\tau^- \rightarrow a_1^- \nu_\tau)$  is equal to  $B(\tau \rightarrow \pi^- \pi^- \pi^+ \nu_\tau \text{ (ex. } K^0, \omega)) + B(\tau \rightarrow \pi^- 2\pi^0 \nu_\tau \text{ (ex. } K^0))$ .

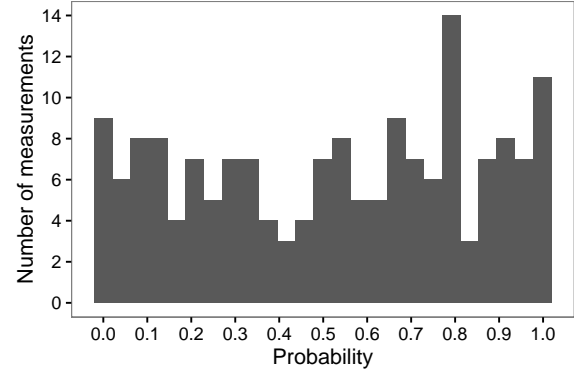


**Figure 1:** Pulls of individual measurements against the respective fitted quantity. No scale factor is used.

In some cases, constraints describe approximate relations that nevertheless hold within the present experimental precision. For instance, the constraint  $B(\tau \rightarrow K^- K^- K^+ \nu_\tau) = B(\tau \rightarrow K^- \phi \nu_\tau) \times B(\phi \rightarrow K^+ K^-)$  is justified within the current experimental evidence.

In the fit, scale factors are applied to the published uncertainties of measurements only if significant inconsistency between different measurements remain after accounting for all relevant uncertainties and correlations. After examining the data and the fit pulls, it has been decided to apply just one scale factor of 5.4 on the measurements of  $B(\tau \rightarrow K^- K^- K^+ \nu_\tau)$ . The scale factor has been computed and applied according to the standard PDG procedure. Without the scale factor applied, the  $\chi^2$  probability of the fit is about 2%. On a per-measurement basis, the pull distribution in figure 1 indicates that just a few measurements have more than  $3\sigma$  pulls. (The uncertainties to obtain the pulls are computed using the measurements variance matrix and the variance matrix of the result, accounting for the fact that the variance matrix of the result is obtained from the measurement variance with the fit.) The pull probability distribution in figure 2 is reasonably flat. With many measurements some entries on the tails of the normal distribution must be expected. There are 169 pulls, one per measurement. They are partially correlated, and the effective number of independent pulls is equal to the number of degrees of freedom of the fit, 124. Only the  $\tau \rightarrow K^- K^- K^+ \nu_\tau$  decay mode has a pull that is inconsistent at the level of more than  $3\sigma$  even if considered as the largest pull in a set of 124. This confirms the choice of adopting just that one scale factor.

After scaling the error the 2016 constrained fit has a  $\chi^2$  of 134.9 for 124 degrees of freedom, corresponding to a  $\chi^2$  probability of 24%. We use 169 measurements and 84 constraints on the branching fractions and ratios to determine 129 quantities, consisting of 112 branching fractions and 17 branching ratios. A total of 85 quantities have at least one measurement in the fit. The constraints include the unitarity constraint on the sum of all the exclusive  $\tau$  decay modes,  $B_{\text{all}} = 1$ . If the unitarity



**Figure 2:** Probability of individual measurement pulls against the respective fitted quantity. No scale factor is used.

constraint is released, the fit result for  $B_{\text{all}}$  is consistent with unitarity with  $1 - B_{\text{all}} = (0.07 \pm 0.10)\%$ .

For the convenience of summarizing the fit results, we list in the following the values and uncertainties for a set of 46 “basis” decay modes, from which all remaining branching fractions and ratios can be obtained using the constraints. Unlike in previous editions, the basis decay modes are not intended to sum up to 1. The new unitarity constraint corresponds to a linear combination of the basis modes weighted by the coefficients listed in the following. The corresponding correlation matrix is listed in the  $\tau$  Listings.

decay mode	fit result (%)	coefficient
$\mu^- \bar{\nu}_\mu \nu_\tau$	$17.3936 \pm 0.0384$	1.0000
$e^- \bar{\nu}_e \nu_\tau$	$17.8174 \pm 0.0399$	1.0000
$\pi^- \nu_\tau$	$10.8165 \pm 0.0512$	1.0000
$K^- \nu_\tau$	$0.6964 \pm 0.0096$	1.0000
$\pi^- \pi^0 \nu_\tau$	$25.4940 \pm 0.0893$	1.0000
$K^- \pi^0 \nu_\tau$	$0.4329 \pm 0.0148$	1.0000
$\pi^- 2\pi^0 \nu_\tau$ (ex. $K^0$ )	$9.2595 \pm 0.0964$	1.0021
$K^- 2\pi^0 \nu_\tau$ (ex. $K^0$ )	$0.0648 \pm 0.0218$	1.0000
$\pi^- 3\pi^0 \nu_\tau$ (ex. $K^0$ )	$1.0428 \pm 0.0707$	1.0000
$K^- 3\pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	$0.0478 \pm 0.0212$	1.0000
$h^- 4\pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	$0.1119 \pm 0.0391$	1.0000
$\pi^- \bar{K}^0 \nu_\tau$	$0.8395 \pm 0.0140$	1.0000
$K^- \bar{K}^0 \nu_\tau$	$0.1479 \pm 0.0053$	1.0000
$\pi^- \bar{K}^0 \pi^0 \nu_\tau$	$0.3821 \pm 0.0129$	1.0000
$K^- \pi^0 K^0 \nu_\tau$	$0.1503 \pm 0.0071$	1.0000
$\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau$ (ex. $K^0$ )	$0.0263 \pm 0.0226$	1.0000
$\pi^- K_S^0 K_S^0 \nu_\tau$	$0.0233 \pm 0.0007$	2.0000
$\pi^- K_S^0 K_L^0 \nu_\tau$	$0.1080 \pm 0.0241$	1.0000
$\pi^- \pi^0 K_S^0 K_S^0 \nu_\tau$	$0.0018 \pm 0.0002$	2.0000
$\pi^- \pi^0 K_S^0 K_L^0 \nu_\tau$	$0.0325 \pm 0.0119$	1.0000
$\bar{K}^0 h^- h^- h^+ \nu_\tau$	$0.0247 \pm 0.0199$	1.0000
$\pi^- \pi^- \pi^+ \nu_\tau$ (ex. $K^0, \omega$ )	$8.9870 \pm 0.0514$	1.0021
$\pi^- \pi^- \pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	$2.7404 \pm 0.0710$	1.0000
$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	$0.0980 \pm 0.0356$	1.0000

$\pi^- K^- K^+ \nu_\tau$	$0.1435 \pm 0.0027$	1.0000
$\pi^- K^- K^+ \pi^0 \nu_\tau$	$0.0061 \pm 0.0018$	1.0000
$\pi^- \pi^0 \eta \nu_\tau$	$0.1389 \pm 0.0072$	1.0000
$K^- \eta \nu_\tau$	$0.0155 \pm 0.0008$	1.0000
$K^- \pi^0 \eta \nu_\tau$	$0.0048 \pm 0.0012$	1.0000
$\pi^- \bar{K}^0 \eta \nu_\tau$	$0.0094 \pm 0.0015$	1.0000
$\pi^- \pi^+ \pi^- \eta \nu_\tau$ (ex. $K^0$ )	$0.0219 \pm 0.0013$	1.0000
$K^- \omega \nu_\tau$	$0.0410 \pm 0.0092$	1.0000
$h^- \pi^0 \omega \nu_\tau$	$0.4085 \pm 0.0419$	1.0000
$K^- \phi \nu_\tau$	$0.0044 \pm 0.0016$	0.8310
$\pi^- \omega \nu_\tau$	$1.9494 \pm 0.0645$	1.0000
$K^- \pi^- \pi^+ \nu_\tau$ (ex. $K^0, \omega$ )	$0.2927 \pm 0.0068$	1.0000
$K^- \pi^- \pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	$0.0394 \pm 0.0142$	1.0000
$\pi^- 2\pi^0 \omega \nu_\tau$ (ex. $K^0$ )	$0.0071 \pm 0.0016$	1.0000
$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1$ )	$0.0014 \pm 0.0027$	1.0000
$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega, f_1$ )	$0.0769 \pm 0.0030$	1.0000
$K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0$ )	$0.0001 \pm 0.0001$	1.0000
$2\pi^- \pi^+ \omega \nu_\tau$ (ex. $K^0$ )	$0.0084 \pm 0.0006$	1.0000
$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1$ )	$0.0038 \pm 0.0009$	1.0000
$K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$0.0001 \pm 0.0001$	1.0000
$\pi^- f_1 \nu_\tau$ ( $f_1 \rightarrow 2\pi^- 2\pi^+$ )	$0.0052 \pm 0.0004$	1.0000
$\pi^- 2\pi^0 \eta \nu_\tau$	$0.0194 \pm 0.0038$	1.0000

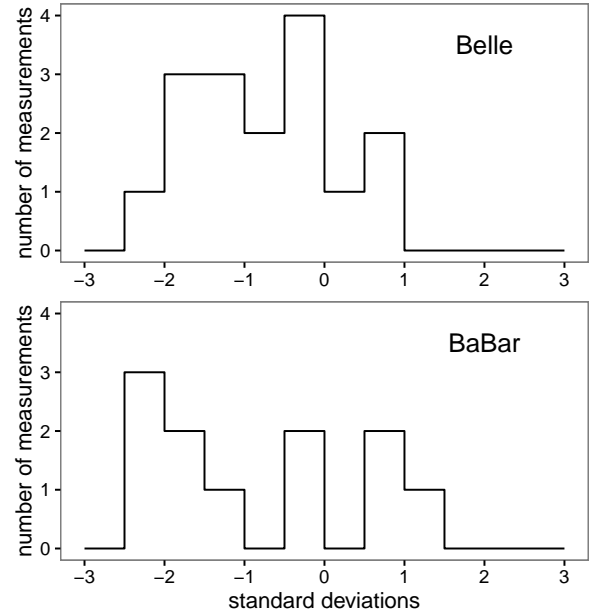
Applying the fit procedure on the PDG 2015 inputs, the fit results differ from the 2015 fit by at most 20% of their uncertainty, for fitted quantities that have measurements with asymmetric errors, and by at most 5% of their uncertainty for the other quantities. The differences originate from the different treatment of asymmetric errors. The present fit procedure symmetrizes the errors as  $\sigma_{\text{symm}}^2 = (\sigma_+^2 + \sigma_-^2)/2$ , while the PDG 2015 fit did model the asymmetric error distributions in the fit. Comparing the results of the previous edition with the current fit, there are differences up to 2.3 times the fitted quantity uncertainty ( $2.3\sigma$ ) for quantities that have no measurement included in the fit and are derived through the constraints. Those differences arise mainly from three changes: the unitarity constraint has been updated to accommodate several additional decay modes, the definitions of the respective quantities have been updated to use the additional decay modes, and the parameters of all constraints (typically,  $K, \eta, \omega$  branching fractions) have been updated to the values reported in the last published PDG edition. For quantities that have measurements in the fit, the fitted values changed at most by  $1.1\sigma$ , reflecting the inclusion of several additional measurements, especially on high-multiplicity decay modes. The uncertainties on the fit results are generally smaller than in 2015 because only one error scale factor is used and some additional measurements have been used.

In defining the fit constraints and in selecting the modes that sum up to one we made some assumptions and choices. We assume that some channels, like  $\tau^- \rightarrow \pi^- K^+ \pi^- \geq 0\pi^0 \nu_\tau$  and  $\tau^- \rightarrow \pi^+ K^- K^- \geq 0\pi^0 \nu_\tau$ , have negligible branching fractions as expected from the Standard Model, even if the experimental limits for these branching fractions are not very stringent. The 95% confidence level upper limits are  $B(\tau^- \rightarrow \pi^- K^+ \pi^- \geq$

$0\pi^0 \nu_\tau) < 0.25\%$  and  $B(\tau^- \rightarrow \pi^+ K^- K^- \geq 0\pi^0 \nu_\tau) < 0.09\%$ , values not so different from measured branching fractions for allowed 3-prong modes containing charged kaons. For decays to final states containing one neutral kaon we assume that the branching fraction with the  $K_L^0$  are the same as the corresponding one with a  $K_S^0$ . On decays with two neutral kaons we assume that the branching fractions with  $K_L^0 K_L^0$  are the same as the ones with  $K_S^0 K_S^0$ .

#### **BaBar and Belle measure on average lower branching fractions and ratios.**

We compare the BaBar and Belle measurements with the results of a fit where all their measurements have been excluded. We find that that BaBar and Belle tend to measure lower  $\tau$  branching fractions and ratios than the other experiments. Figure 3 shows histograms of the 27 normalized differences between the  $B$ -factory measurements and the respective non- $B$ -factory fit results. The normalization is the uncertainty on the difference. The average normalized difference between the two sets of measurements is  $-0.8\sigma$  ( $-0.8\sigma$  for the 16 Belle measurements and  $-0.9\sigma$  for the 11 BaBar measurements).



**Figure 3:** Distribution of the normalized difference between the 27  $B$ -factory measurements and non- $B$ -factory measurements. The list includes 16 measurements of branching fractions and ratios published by the Belle collaboration and 11 by the BaBar collaboration that are used in the fit and for which non- $B$ -factory measurements exist.

#### **Overconsistency of Leptonic Branching Fraction Measurements.**

As observed in the previous editions of this review, measurements of the leptonic branching fractions are more consistent with each other than expected from the quoted errors on the individual measurements. The  $\chi^2$  is 0.34 for  $B_e$  and 0.08 for

# Lepton Particle Listings

$\tau$

$B_\mu$ . Assuming normal errors, the probability of a smaller  $\chi^2$  is 1.3% for  $B_e$  and 0.08% for  $B_\mu$ .

## Technical implementation of the fit.

The fit computes a set of quantities denoted with  $q_i$  by minimizing a  $\chi^2$  while respecting a series of equality constraints on the  $q_i$ . The  $\chi^2$  is computed using the measurements  $m_i$  and their covariance matrix  $E_{ij}$  as  $\chi^2 = (m_i - A_{ik}q_k)^t E_{ij}^{-1} (m_j - A_{jl}q_l)$  where the model matrix  $A_{ij}$  is used to get the vector of the predicted measurements  $m'_i$  from the vector of the fit parameters  $q_j$  as  $m'_i = A_{ij}q_j$ . In this particular implementation the measurements are grouped by the quantity that they measure, and all quantities with at least one measurement correspond to a fit parameter. Therefore, the matrix  $A_{ij}$  has one row per measurement  $m_i$  and one column per fitted quantity  $q_j$ , with unity coefficients for the rows and column that identify a measurement  $m_i$  of the quantity  $q_j$ , respectively. The constraints are equations involving the fit parameters. The fit does not impose limitations on the functional form of the constraints. In summary, the fit requires:

$$\min(m_i - A_{ik}q_k)^t E_{ij}^{-1} (m_j - A_{jl}q_l), \quad (1a)$$

$$\text{subjected to } f_r(q_s) - c_r = 0, \quad (1b)$$

where the left term of Eq. (1b) defines the constraint expressions. Using the method of Lagrange multipliers, a set of equations is obtained by taking the derivatives with respect to the fitted quantities  $q_k$  and the Lagrange multipliers  $\lambda_r$  of the sum of the  $\chi^2$  and the constraint expressions multiplied by the Lagrange multipliers  $\lambda_r$ , one for each constraint:

$$\min \left[ (A_{ik}q_k - m_i)^t E_{ij}^{-1} (A_{jl}q_l - m_j) + 2\lambda_r (f_r(q_s) - c_r) \right] \quad (2a)$$

$$(\partial/\partial q_k, \partial/\partial \lambda_r)[\text{expression above}] = 0 \quad (2b)$$

Eq. (2b) defines a set of equations for the vector of the unknowns  $(q_k, \lambda_r)$ , some of which may be non-linear, in case of non-linear constraints. An iterative minimization procedure approximates at each step the non-linear constraint expressions by their first order Taylor expansion around the current values of the fitted quantities,  $\bar{q}_s$ :

$$f_r(q_s) - c_r = f_r(\bar{q}_s) + \frac{\partial f_r(q_s)}{\partial q_s} \bigg|_{\bar{q}_s} (q_s - \bar{q}_s) - c_r, \quad (3a)$$

which can be written as

$$B_{rs}q_s - c'_r, \quad (3b)$$

where  $c'_r$  are the resulting constant known terms, independent of  $q_s$  at first order. After linearization, the differentiation by  $q_k$  and  $\lambda_r$  is trivial and leads to a set of linear equations

$$A_{ki}^t E_{ij}^{-1} A_{jl} q_l + B_{kr}^t \lambda_r = A_{ki}^t E_{ij}^{-1} m_j \quad (4a)$$

$$B_{rs}q_s = c'_r, \quad (4b)$$

which can be expressed as:

$$F_{ij} u_j = v_i \quad (5)$$

where  $u_j = (q_k, \lambda_r)$  and  $v_i$  is the vector of the known constant terms running over the index  $k$  and then  $r$  in the right terms of Eq. (4a) and Eq. (4b), respectively. Solving the equation set in Eq. (5) by matrix inversion gives the the fitted quantities and their variance and covariance matrix, using the measurements and their variance and covariance matrix. The fit procedure starts by computing the linear approximation of the non-linear constraint expressions around the quantities seed values. With an iterative procedure, the unknowns are updated at each step by solving the equations and the equations are then linearized around the updated values, until the variation of the fitted unknowns is reduced below a numerically small threshold.

## References

1. D. Asner *et al.* (HFAG), [arXiv:1010.1589](#); Y. Amhis *et al.* (HFAG), [arXiv:1412.7515](#).
2. See the Introduction section of this edition of the *Review of Particle Physics*.
3. S. Schael *et al.* (ALEPH Collab.), Phys. Reports **421**, 191 (2005).
4. J. Abdallah *et al.* (DELPHI Collab.), Eur. Phys. J. **C46**, 1 (2006).
5. J.P. Lees *et al.* (BABAR Collab.), Phys. Rev. **D86**, 092010 (2012), [[arXiv:1209.2734](#)].
6. J.P. Lees *et al.* (BABAR Collab.), Phys. Rev. **D86**, 092013 (2012), [[arXiv:1208.0376](#)].
7. S. Ryu *et al.* (BELLE Collab.), Phys. Rev. **D89**, 072009 (2014).
8. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **100**, 011801 (2008).
9. K. Inami *et al.* (BELLE Collab.), Phys. Lett. **B643**, 5 (2006).
10. M. Zielinski *et al.*, Phys. Rev. Lett. **52**, 1195 (1984).

$(\Gamma(\tau^+) - \Gamma(\tau^-)) / (\Gamma(\tau^+) + \Gamma(\tau^-))$			
$\tau^\pm \rightarrow \pi^\pm K_S^0 \nu_\tau$ (RATE DIFFERENCE) / (RATE SUM)			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>-0.36 \pm 0.23 \pm 0.11</math></b>	LEES	12M BABR	476 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

## $\tau^-$ BRANCHING RATIOS

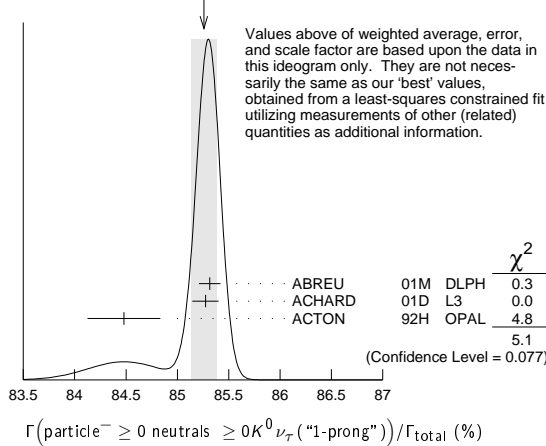
$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")}) / \Gamma_{\text{total}}$	$\Gamma_1 / \Gamma$
$\Gamma_1 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + \Gamma_{36} + \Gamma_{38} + \Gamma_{41} + \Gamma_{43} + \Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{50} + \Gamma_{52} + \Gamma_{56} + \Gamma_{57} + 0.7212\Gamma_{148} + 0.7212\Gamma_{150} + 0.7212\Gamma_{152} + 0.7212\Gamma_{154} + 0.342\Gamma_{168} + 0.0828\Gamma_{176} + 0.0828\Gamma_{177} + 0.0828\Gamma_{178}) / \Gamma$	

The charged particle here can be  $e$ ,  $\mu$ , or hadron. In many analyses, the sum of the topological branching fractions (1, 3, and 5 prongs) is constrained to be unity. Since the 5-prong fraction is very small, the measured 1-prong and 3-prong fractions are highly correlated and cannot be treated as independent quantities in our overall fit. We arbitrarily choose to use the 3-prong fraction in our fit, and leave the 1-prong fraction out. We do, however, use these 1-prong measurements in our average below. The measurements used only for the average are marked "avg," whereas "f&a" marks a result used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
85.24 ±0.06	OUR FIT			
85.26 ±0.13	OUR AVERAGE	Error includes scale factor of 1.6. See the ideogram below.		
● ● ● We use the following data for averages but not for fits. ● ● ●				
85.316 ±0.093 ±0.049	78k	<sup>1</sup> ABREU	01M DLPH	1992-1995 LEP runs
85.274 ±0.105 ±0.073		<sup>2</sup> ACHARD	01D L3	1992-1995 LEP runs
84.48 ±0.27 ±0.23		ACTON	92H OPAL	1990-1991 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
85.45 +0.69 -0.73 ±0.65		DECAMP	92c ALEP	Repl. by SCHAEEL 05c

- <sup>1</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow 3\text{-prong})$  and  $B(\tau \rightarrow 5\text{-prong})$  are  $-0.98$  and  $-0.08$  respectively.
- <sup>2</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow "3\text{-prong}")$  and  $B(\tau \rightarrow "5\text{-prong}")$  are  $-0.978$  and  $-0.082$  respectively.

WEIGHTED AVERAGE  
85.26±0.13 (Error scaled by 1.6)



$\Gamma(\text{partic}^- \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_2 / \Gamma$

$\Gamma_2 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.6534\Gamma_{36} + 0.6534\Gamma_{38} + 0.6534\Gamma_{41} + 0.6534\Gamma_{43} + 0.6534\Gamma_{45} + 0.0942\Gamma_{48} + 0.3069\Gamma_{49} + \Gamma_{50} + 0.0942\Gamma_{52} + 0.3069\Gamma_{56} + \Gamma_{57} + 0.7212\Gamma_{148} + 0.7212\Gamma_{150} + 0.7212\Gamma_{152} + 0.4712\Gamma_{154} + 0.1049\Gamma_{168} + 0.0828\Gamma_{176} + 0.0828\Gamma_{177} + 0.0828\Gamma_{178}) / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>84.58 ± 0.06 OUR FIT</b>				
<b>85.1 ± 0.4 OUR AVERAGE</b>				
• • • We use the following data for averages but not for fits. • • •				
85.6 ± 0.6 ± 0.3	3300	<sup>1</sup> ADEVA 91F L3	$E_{\text{cm}}^{\text{ee}} = 88.3\text{--}94.3 \text{ GeV}$	
84.9 ± 0.4 ± 0.3		BEHREND 89B CELL	$E_{\text{cm}}^{\text{ee}} = 14\text{--}47 \text{ GeV}$	
84.7 ± 0.8 ± 0.6		<sup>2</sup> AIHARA 87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
86.4 ± 0.3 ± 0.3		ABACHI 89B HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
87.1 ± 1.0 ± 0.7		<sup>3</sup> BURCHAT 87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
87.2 ± 0.5 ± 0.8		SCHMIDKE 86 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
84.7 ± 1.1 $^{+1.6}_{-1.3}$	169	<sup>4</sup> ALTHOFF 85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$	
86.1 ± 0.5 ± 0.9		BARTEL 85F JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
87.8 ± 1.3 ± 0.9		<sup>5</sup> BERGER 85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
86.7 ± 0.3 ± 0.6		FERNANDEZ 85 MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	

- <sup>1</sup> Not independent of ADEVA 91F  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  value.
- <sup>2</sup> Not independent of AIHARA 87B  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  values.
- <sup>3</sup> Not independent of SCHMIDKE 86 value (also not independent of BURCHAT 87 value for  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ ).
- <sup>4</sup> Not independent of ALTHOFF 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  values.
- <sup>5</sup> Not independent of (1-prong +  $0\pi^0$ ) and (1-prong +  $\geq 1\pi^0$ ) values.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_3 / \Gamma$

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.39 ± 0.04 OUR FIT</b>				
<b>17.33 ± 0.05 OUR AVERAGE</b>				
17.319 ± 0.070 ± 0.032	54k	<sup>1</sup> SCHAEEL 05c ALEP	1991-1995 LEP runs	
17.34 ± 0.09 ± 0.06	31.4k	ABBIENDI 03 OPAL	1990-1995 LEP runs	
17.342 ± 0.110 ± 0.067	21.5k	<sup>2</sup> ACCIARRI 01F L3	1991-1995 LEP runs	
17.325 ± 0.095 ± 0.077	27.7k	ABREU 99x DLPH	1991-1995 LEP runs	
• • • We use the following data for averages but not for fits. • • •				
17.37 ± 0.08 ± 0.18		<sup>3</sup> ANASTASSOV 97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17.31 ± 0.11 ± 0.05	20.7k	BUSKULIC 96c ALEP	Repl. by SCHAEEL 05c	
17.02 ± 0.19 ± 0.24	6586	ABREU 95T DLPH	Repl. by ABREU 99x	
17.36 ± 0.27	7941	AKERS 95i OPAL	Repl. by ABBIENDI 03	
17.6 ± 0.4 ± 0.4	2148	ADRIANI 93M L3	Repl. by ACCIARRI 01F	
17.4 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT 93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
17.35 ± 0.41 ± 0.37		DECAMP 92c ALEP	1989-1990 LEP runs	
17.7 ± 0.8 ± 0.4	568	BEHREND 90 CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$	
17.4 ± 1.0	2197	ADEVA 88 MRKJ	$E_{\text{cm}}^{\text{ee}} = 14\text{--}16 \text{ GeV}$	
17.7 ± 1.2 ± 0.7		AIHARA 87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	

18.3 ± 0.9 ± 0.8		BURCHAT 87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.6 ± 0.8 ± 0.7	558	<sup>5</sup> BARTEL 86D JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
12.9 ± 1.7 $^{+0.7}_{-0.5}$		ALTHOFF 85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$
18.0 ± 0.9 ± 0.5	473	<sup>5</sup> ASH 85B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.0 ± 1.0 ± 0.6		<sup>6</sup> BALTRUSAITIS...85 MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$
19.4 ± 1.6 ± 1.7	153	BERGER 85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
17.6 ± 2.6 ± 2.1	47	BEHREND 83c CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$
17.8 ± 2.0 ± 1.8		BERGER 81B PLUT	$E_{\text{cm}}^{\text{ee}} = 9\text{--}32 \text{ GeV}$

- <sup>1</sup> See footnote to SCHAEEL 05c  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.
- <sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  is 0.08.
- <sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(e^- \bar{\nu}_e \nu_\tau)$ ,  $B(\mu^- \bar{\nu}_\mu \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau) / B(e^- \bar{\nu}_e \nu_\tau)$  are 0.50, 0.58, 0.50, and 0.08 respectively.
- <sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}^2$  values.
- <sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau) / B("1 \text{ prong} ")$  and  $B("1 \text{ prong} ") = 0.855$ .
- <sup>6</sup> Error correlated with BALTRUSAITIS 85  $e\nu\bar{\nu}$  value.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma_{\text{total}}$	VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.368 ± 0.010 OUR AVERAGE</b>					
0.369 ± 0.003 ± 0.010	16k	<sup>1</sup> LEES 15G BABR	431 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
0.361 ± 0.016 ± 0.035		<sup>2</sup> BERGFELD 00 CLEO		$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.30 ± 0.04 ± 0.05	116	<sup>3</sup> ALEXANDER 96s OPAL	1991-1994 LEP runs		
0.23 ± 0.10	10	<sup>4</sup> WU 90 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$		

- <sup>1</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .
- <sup>2</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ . For  $E_\gamma^* > 20 \text{ MeV}$ , they quote  $(3.04 \pm 0.14 \pm 0.30) \times 10^{-3}$ .
- <sup>3</sup> ALEXANDER 96s impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma > 20 \text{ MeV}$ .
- <sup>4</sup> WU 90 reports  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) = 0.013 \pm 0.006$ , which is converted to  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma_{\text{total}}$  using  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma) / \Gamma_{\text{total}} = 17.35\%$ . Requirements on detected  $\gamma$ 's correspond to a  $\tau$  rest frame energy cutoff  $E_\gamma > 37 \text{ MeV}$ .

$\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_5 / \Gamma$

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.82 ± 0.04 OUR FIT</b>				
<b>17.82 ± 0.05 OUR AVERAGE</b>				
17.837 ± 0.072 ± 0.036	56k	<sup>1</sup> SCHAEEL 05c ALEP	1991-1995 LEP runs	
17.806 ± 0.104 ± 0.076	24.7k	<sup>2</sup> ACCIARRI 01F L3	1991-1995 LEP runs	
17.81 ± 0.09 ± 0.06	33.1k	ABBIENDI 99H OPAL	1991-1995 LEP runs	
17.877 ± 0.109 ± 0.110	23.3k	ABREU 99x DLPH	1991-1995 LEP runs	
17.76 ± 0.06 ± 0.17		<sup>3</sup> ANASTASSOV 97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17.78 ± 0.10 ± 0.09	25.3k	ALEXANDER 96D OPAL	Repl. by ABBI- ENDI 99H	
17.79 ± 0.12 ± 0.06	20.6k	BUSKULIC 96c ALEP	Repl. by SCHAEEL 05c	
17.51 ± 0.23 ± 0.31	5059	ABREU 95T DLPH	Repl. by ABREU 99x	
17.9 ± 0.4 ± 0.4	2892	ADRIANI 93M L3	Repl. by ACCIARRI 01F	
17.5 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT 93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
17.97 ± 0.14 ± 0.23	3970	AKERIB 92 CLEO	Repl. by ANAS- TASSOV 97	
19.1 ± 0.4 ± 0.6	2960	<sup>5</sup> AMMAR 92 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5\text{--}10.9 \text{ GeV}$	
18.09 ± 0.45 ± 0.45		DECAMP 92c ALEP	Repl. by SCHAEEL 05c	
17.0 ± 0.5 ± 0.6	1.7k	ABACHI 90 HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
18.4 ± 0.8 ± 0.4	644	BEHREND 90 CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$	
16.3 ± 0.3 ± 3.2		JANSSEN 89 CBAL	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
18.4 ± 1.2 ± 1.0		AIHARA 87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
19.1 ± 0.8 ± 1.1		BURCHAT 87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
16.8 ± 0.7 ± 0.9	515	<sup>5</sup> BARTEL 86D JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
20.4 ± 3.0 $^{+1.4}_{-0.9}$		ALTHOFF 85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$	
17.8 ± 0.9 ± 0.6	390	<sup>5</sup> ASH 85B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
18.2 ± 0.7 ± 0.5		<sup>6</sup> BALTRUSAITIS...85 MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$	
13.0 ± 1.9 ± 2.9		BERGER 85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
18.3 ± 2.4 ± 1.9	60	BEHREND 83c CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$	
16.0 ± 1.3	459	<sup>7</sup> BACINO 78B DLCO	$E_{\text{cm}}^{\text{ee}} = 3.1\text{--}7.4 \text{ GeV}$	

- <sup>1</sup> Correlation matrix for SCHAEEL 05c branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$
- (5)  $\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$

## Lepton Particle Listings

 $\tau$ 

- (6)  $\Gamma(\tau^- \rightarrow \pi^- 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (7)  $\Gamma(\tau^- \rightarrow h^- 4\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}}$   
 (8)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$   
 (9)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (10)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (11)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 3\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   
 (12)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (13)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)	(12)
(2)	-20										
(3)	-9	-6									
(4)	-16	-12	2								
(5)	-5	-5	-17	-37							
(6)	0	-4	-15	2	-27						
(7)	-2	-4	-24	-15	20	-47					
(8)	-14	-9	15	-5	-17	-14	-8				
(9)	-13	-12	-25	-30	4	-2	16	-15			
(10)	0	-2	-23	-14	4	10	13	-6	-17		
(11)	1	0	-5	1	4	6	0	-9	-2	-11	
(12)	0	1	9	4	-8	-4	-6	9	-5	-4	-2
(13)	1	-4	-3	-5	3	2	-4	-3	-1	4	1

<sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  is 0.08.

<sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.50, -0.42, 0.48, and -0.39 respectively.

<sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  values.

<sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

<sup>6</sup> Error correlated with BALTRUSAITIS 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ .

<sup>7</sup> BACINO 78B value comes from fit to events with  $e^\pm$  and one other nonelectron charged prong.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$					$\Gamma_3/\Gamma_5$	
Standard Model prediction including mass effects is 0.9726.						
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>97.62 ± 0.28 OUR FIT</b>						
<b>97.9 ± 0.4 OUR AVERAGE</b>						
97.96 ± 0.16 ± 0.36	731k	<sup>1</sup> AUBERT	10F	BABR 467 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV		
97.77 ± 0.63 ± 0.87		<sup>2</sup> ANASTASSOV 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV		
99.7 ± 3.5 ± 4.0		ALBRECHT 92D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV		

<sup>1</sup> Correlation matrix for AUBERT 10F branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$   
 (2)  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$   
 (3)  $\Gamma(\tau^- \rightarrow K^- \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$

(1)	(2)
(2)	0.25
(3)	0.12
	0.33

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(e \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.58, -0.42, 0.07, and 0.45 respectively.

$\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$	
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>1.84 ± 0.05 OUR AVERAGE</b>						
1.847 ± 0.015 ± 0.052	18k	<sup>1</sup> LEES	15G	BABR 431 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV		
1.75 ± 0.06 ± 0.17		<sup>2</sup> BERGFELD	00	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV		

<sup>1</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

<sup>2</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

$\Gamma(h^- \geq 0K_L^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
$\Gamma_7/\Gamma = (\Gamma_9 + \Gamma_{10} + \frac{1}{2}\Gamma_{36} + \frac{1}{2}\Gamma_{38} + \Gamma_{50})/\Gamma$					
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>12.03±0.05 OUR FIT</b>					
<b>12.2 ±0.4 OUR AVERAGE</b>					
12.47 ±0.26 ±0.43	2967	<sup>1</sup> ACCIARRI	95	L3	1992 LEP run
12.4 ±0.7 ±0.7	283	<sup>2</sup> ABREU	92N	DLPH	1990 LEP run
12.1 ±0.7 ±0.5	309	ALEXANDER	91D	OPAL	1990 LEP run
• • • We use the following data for averages but not for fits. • • •					
11.3 ±0.5 ±0.8	798	<sup>3</sup> FORD	87	MAC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
12.44 ±0.11 ±0.11	15k	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by SCHAE 05c
11.7 ±0.6 ±0.8		<sup>5</sup> ALBRECHT	92D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV
12.98 ±0.44 ±0.33		<sup>6</sup> DECAMP	92c	ALEP	Repl. by SCHAE 05c
12.3 ±0.9 ±0.5	1338	BEHREND	90	CELL	$E_{\text{cm}}^{\text{ee}} = 35$ GeV

11.1 ± 1.1 ± 1.4		<sup>7</sup> BURCHAT	87	MRK2 $E_{\text{cm}}^{\text{ee}} = 29$ GeV
12.3 ± 0.6 ± 1.1	328	<sup>8</sup> BARTEL	86D	JADE $E_{\text{cm}}^{\text{ee}} = 34.6$ GeV
13.0 ± 2.0 ± 4.0		BERGER	85	PLUT $E_{\text{cm}}^{\text{ee}} = 34.6$ GeV
11.2 ± 1.7 ± 1.2	34	<sup>9</sup> BEHREND	83c	CELL $E_{\text{cm}}^{\text{ee}} = 34$ GeV

<sup>1</sup> ACCIARRI 95 with 0.65% added to remove their correction for  $\pi^- K_L^0$  backgrounds.

<sup>2</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>3</sup> FORD 87 result for  $B(\pi^- \nu_\tau)$  with 0.67% added to remove their  $K^-$  correction and adjusted for 1992 B("1 prong").

<sup>4</sup> BUSKULIC 96 quote  $11.78 \pm 0.11 \pm 0.13$  We add 0.66 to undo their correction for unseen  $K_L^0$  and modify the systematic error accordingly.

<sup>5</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ ,  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)$ , and  $\Gamma(h^- \geq 0K_L^0 \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  values.

<sup>6</sup> DECAMP 92c quote  $B(h^- \geq 0K_L^0 \geq 0(K_S^0 \rightarrow \pi^+ \pi^-) \nu_\tau) = 13.32 \pm 0.44 \pm 0.33$ . We subtract 0.35 to correct for their inclusion of the  $K_S^0$  decays.

<sup>7</sup> BURCHAT 87 with 1.1% added to remove their correction for  $K^-$  and  $K^*(892)^-$  backgrounds.

<sup>8</sup> BARTEL 86D result for  $B(\pi^- \nu_\tau)$  with 0.59% added to remove their  $K^-$  correction and adjusted for 1992 B("1 prong").

<sup>9</sup> BEHREND 83c quote  $B(\pi^- \nu_\tau) = 9.9 \pm 1.7 \pm 1.3$  after subtracting  $1.3 \pm 0.5$  to correct for  $B(K^- \nu_\tau)$ .

$\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma = (\Gamma_9 + \Gamma_{10})/\Gamma$	
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>11.51 ± 0.05 OUR FIT</b>					
<b>11.63 ± 0.12 OUR AVERAGE</b>					
11.571 ± 0.120 ± 0.114	19k	<sup>1</sup> ABDALLAH	06A	DLPH 1992-1995 LEP runs	
11.98 ± 0.13 ± 0.16		ACKERSTAFF 98M	OPAL	1991-1995 LEP runs	
11.52 ± 0.05 ± 0.12		<sup>2</sup> ANASTASSOV 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

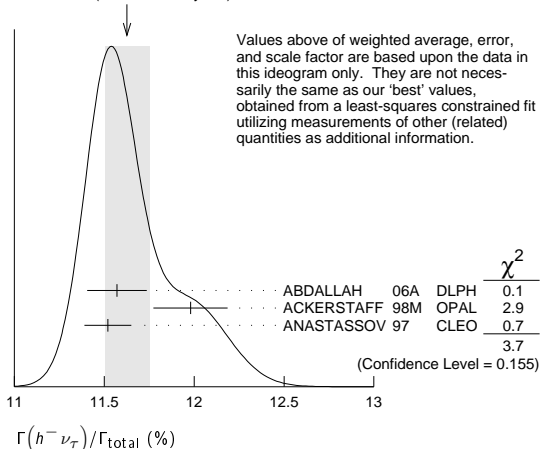
<sup>1</sup> Correlation matrix for ABDALLAH 06A branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$   
 (2)  $\Gamma(\tau^- \rightarrow h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   
 (3)  $\Gamma(\tau^- \rightarrow h^- \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (4)  $\Gamma(\tau^- \rightarrow h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (5)  $\Gamma(\tau^- \rightarrow h^- \geq 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (6)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (7)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (8)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (9)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (10)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   
 (11)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
(2)	-34								
(3)	-47	56							
(4)	6	-66	15						
(5)	-6	38	11	-86					
(6)	-7	-8	15	0	-2				
(7)	-2	-1	-5	-3	3	-53			
(8)	-4	-4	-13	-4	-2	-56	75		
(9)	-1	-1	-4	3	-6	26	-78	-16	
(10)	-1	-1	1	0	0	-2	-3	-1	3
(11)	0	0	0	0	0	1	0	-5	5

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(e \bar{\nu}_e \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.50, 0.48, 0.07, and 0.63 respectively.

WEIGHTED AVERAGE  
 11.63 ± 0.12 (Error scaled by 1.4)



$\Gamma(h^-\nu_\tau)/\Gamma(e^-\nu_e\nu_\tau)$   $\Gamma_8/\Gamma_5 = (\Gamma_9 + \Gamma_{10})/\Gamma_5$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>64.62 ± 0.33 OUR FIT</b>				
<b>64.0 ± 0.7 OUR AVERAGE</b>				Error includes scale factor of 1.6.
• • •				We use the following data for averages but not for fits. • • •
63.33 ± 0.14 ± 0.61	394k	<sup>1</sup> AUBERT	10F BABR	467 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
64.84 ± 0.41 ± 0.60		<sup>2</sup> ANASTASSOV	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV
<sup>1</sup> Not independent of AUBERT 10F $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \nu_e \nu_\tau)$ and $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \nu_e \nu_\tau)$ .				
<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of $B(\mu \bar{\nu}_\mu \nu_\tau)$ , $B(e \bar{\nu}_e \nu_\tau)$ , $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ , and $B(h^- \nu_\tau)$ are 0.08, -0.39, 0.45, and 0.63 respectively.				

 $\Gamma(\pi^-\nu_\tau)/\Gamma_{total}$   $\Gamma_9/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.82 ± 0.05 OUR FIT</b>				
<b>10.828 ± 0.070 ± 0.078</b>	38k	<sup>1</sup> SCHAE	05c ALEP	1991-1995 LEP runs
• • •				We do not use the following data for averages, fits, limits, etc. • • •
11.06 ± 0.11 ± 0.14		<sup>2</sup> BUSKULIC	96 ALEP	Repl. by SCHAE 05c
11.7 ± 0.4 ± 1.8	1138	BLOCKER	82D MRK2	$E_{cm}^{ee} = 3.5-6.7$ GeV
<sup>1</sup> See footnote to SCHAE 05c $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{total}$ measurement for correlations with other measurements.				
<sup>2</sup> Not independent of BUSKULIC 96 $B(h^- \nu_\tau)$ and $B(K^- \nu_\tau)$ values.				

 $\Gamma(\pi^-\nu_\tau)/\Gamma(e^-\nu_e\nu_\tau)$   $\Gamma_9/\Gamma_5$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>60.71 ± 0.32 OUR FIT</b>				
<b>59.45 ± 0.14 ± 0.61</b>	369k	<sup>1</sup> AUBERT	10F BABR	467 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
<sup>1</sup> See footnote to AUBERT 10F $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ for correlations with other measurements.				

 $\Gamma(K^-\nu_\tau)/\Gamma_{total}$   $\Gamma_{10}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.696 ± 0.010 OUR FIT</b>				
<b>0.685 ± 0.023 OUR AVERAGE</b>				
0.658 ± 0.027 ± 0.029		<sup>1</sup> ABBIENDI	01J OPAL	1990-1995 LEP runs
0.696 ± 0.025 ± 0.014	2032	BARATE	99K ALEP	1991-1995 LEP runs
0.85 ± 0.18	27	ABREU	94K DLPH	LEP 1992 Z data
0.66 ± 0.07 ± 0.09	99	BATTLE	94 CLEO	$E_{cm}^{ee} \approx 10.6$ GeV
• • •				We do not use the following data for averages, fits, limits, etc. • • •
0.72 ± 0.04 ± 0.04	728	BUSKULIC	96 ALEP	Repl. by BARATE 99K
0.59 ± 0.18	16	MILLS	84 DLCO	$E_{cm}^{ee} = 29$ GeV
1.3 ± 0.5	15	BLOCKER	82B MRK2	$E_{cm}^{ee} = 3.9-6.7$ GeV
<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$ is 0.60.				

 $\Gamma(K^-\nu_\tau)/\Gamma(e^-\nu_e\nu_\tau)$   $\Gamma_{10}/\Gamma_5$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.91 ± 0.05 OUR FIT</b>				
<b>3.882 ± 0.032 ± 0.057</b>	25k	<sup>1</sup> AUBERT	10F BABR	467 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
<sup>1</sup> See footnote to AUBERT 10F $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ for correlations with other measurements.				

 $\Gamma(K^-\nu_\tau)/\Gamma(\pi^-\nu_\tau)$   $\Gamma_{10}/\Gamma_9$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.44 ± 0.09 OUR FIT</b>			
• • •			We use the following data for averages but not for fits. • • •
<b>6.531 ± 0.056 ± 0.093</b>	<sup>1</sup> AUBERT	10F BABR	467 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
<sup>1</sup> Not independent of AUBERT 10F $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \nu_e \nu_\tau)$ and $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \nu_e \nu_\tau)$ .			

 $\Gamma(h^- \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{total}$   $\Gamma_{11}/\Gamma$ 

$\Gamma_{11}/\Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.7212\Gamma_{148} + 0.7212\Gamma_{150} + 0.7212\Gamma_{152} + 0.1107\Gamma_{154} + 0.0828\Gamma_{176} + 0.0828\Gamma_{177} + 0.0828\Gamma_{178})/\Gamma$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>37.00 ± 0.09 OUR FIT</b>				
• • •				We do not use the following data for averages, fits, limits, etc. • • •
36.14 ± 0.33 ± 0.58	<sup>1</sup> AKERS	94E OPAL	1991-1992 LEP runs	
38.4 ± 1.2 ± 1.0	<sup>2</sup> BURCHAT	87 MRK2	$E_{cm}^{ee} = 29$ GeV	
42.7 ± 2.0 ± 2.9	BERGER	85 PLUT	$E_{cm}^{ee} = 34.6$ GeV	
<sup>1</sup> Not independent of ACKERSTAFF 98M $B(h^- \pi^0 \nu_\tau)$ and $B(h^- \geq 2\pi^0 \nu_\tau)$ values.				
<sup>2</sup> BURCHAT 87 quote for $B(\pi^\pm \geq 1 \text{ neutral } \nu_\tau) = 0.378 \pm 0.012 \pm 0.010$ . We add 0.006 to account for contribution from $(K^* - \nu_\tau)$ which they fixed at BR = 0.013.				

 $\Gamma(h^- \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{total}$   $\Gamma_{12}/\Gamma$ 

$\Gamma_{12}/\Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3268\Gamma_{148} + 0.3268\Gamma_{150} + 0.3268\Gamma_{152})/\Gamma$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>36.51 ± 0.09 OUR FIT</b>				
• • •				We use the following data for averages but not for fits. • • •
<b>36.641 ± 0.155 ± 0.127</b>	45k	<sup>1</sup> ABDALLAH	06A DLPH	1992-1995 LEP runs
<sup>1</sup> See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{total}$ measurement for correlations with other measurements.				

 $\Gamma(h^- \pi^0 \nu_\tau)/\Gamma_{total}$   $\Gamma_{13}/\Gamma = (\Gamma_{14} + \Gamma_{16})/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>25.93 ± 0.09 OUR FIT</b>				
<b>25.73 ± 0.16 OUR AVERAGE</b>				
25.67 ± 0.01 ± 0.39	5.4M	FUJIKAWA	08 BELL	72 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
25.740 ± 0.201 ± 0.138	35k	<sup>1</sup> ABDALLAH	06A DLPH	1992-1995 LEP runs
25.89 ± 0.17 ± 0.29		ACKERSTAFF	98M OPAL	1991-1995 LEP runs
25.05 ± 0.35 ± 0.50	6613	ACCIARRI	95 L3	1992 LEP run
25.87 ± 0.12 ± 0.42	51k	<sup>2</sup> ARTUSO	94 CLEO	$E_{cm}^{ee} = 10.6$ GeV
• • •				We do not use the following data for averages, fits, limits, etc. • • •
25.76 ± 0.15 ± 0.13	31k	BUSKULIC	96 ALEP	Repl. by SCHAE 05c
25.98 ± 0.36 ± 0.52		<sup>3</sup> AKERS	94E OPAL	Repl. by ACKERSTAFF 98M
22.9 ± 0.8 ± 1.3	283	<sup>4</sup> ABREU	92N DLPH	$E_{cm}^{ee} = 88.2-94.2$ GeV
23.1 ± 0.4 ± 0.9	1249	<sup>5</sup> ALBRECHT	92Q ARG	$E_{cm}^{ee} = 10$ GeV
25.02 ± 0.64 ± 0.88	1849	DECAMP	92C ALEP	1989-1990 LEP runs
22.0 ± 0.8 ± 1.9	779	ANTREASYAN	91 CBAL	$E_{cm}^{ee} = 9.4-10.6$ GeV
22.6 ± 1.5 ± 0.7	1101	BEHREND	90 CELL	$E_{cm}^{ee} = 35$ GeV
23.1 ± 1.9 ± 1.6		BEHREND	84 CELL	$E_{cm}^{ee} = 14,22$ GeV

- <sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{total}$  measurement for correlations with other measurements.
- <sup>2</sup> ARTUSO 94 reports the combined result from three independent methods, one of which (23% of the  $\tau^- \rightarrow h^- \pi^0 \nu_\tau$ ) is normalized to the inclusive one-prong branching fraction, taken as  $0.854 \pm 0.004$ . Renormalization to the present value causes negligible change.
- <sup>3</sup> AKERS 94e quote  $(26.25 \pm 0.36 \pm 0.52) \times 10^{-2}$ ; we subtract 0.27% from their number to correct for  $\tau^- \rightarrow h^- K_L^0 \nu_\tau$ .
- <sup>4</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^- \rightarrow$  backgrounds.
- <sup>5</sup> ALBRECHT 92Q with 0.5% added to remove their correction for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  background.

 $\Gamma(\pi^- \pi^0 \nu_\tau)/\Gamma_{total}$   $\Gamma_{14}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>25.49 ± 0.09 OUR FIT</b>				
<b>25.46 ± 0.12 OUR AVERAGE</b>				
25.471 ± 0.097 ± 0.085	81k	<sup>1</sup> SCHAE	05c ALEP	1991-1995 LEP runs
• • •				We use the following data for averages but not for fits. • • •
25.36 ± 0.44		<sup>2</sup> ARTUSO	94 CLEO	$E_{cm}^{ee} = 10.6$ GeV
• • •				We do not use the following data for averages, fits, limits, etc. • • •
25.30 ± 0.15 ± 0.13		<sup>3</sup> BUSKULIC	96 ALEP	Repl. by SCHAE 05c
21.5 ± 0.4 ± 1.9	4400	<sup>4,5</sup> ALBRECHT	88L ARG	$E_{cm}^{ee} = 10$ GeV
23.0 ± 1.3 ± 1.7	582	ADLER	87B MRK3	$E_{cm}^{ee} = 3.77$ GeV
25.8 ± 1.7 ± 2.5		<sup>6</sup> BURCHAT	87 MRK2	$E_{cm}^{ee} = 29$ GeV
22.3 ± 0.6 ± 1.4	629	<sup>5</sup> YELTON	86 MRK2	$E_{cm}^{ee} = 29$ GeV

- <sup>1</sup> See footnote to SCHAE 05c  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{total}$  measurement for correlations with other measurements.
- <sup>2</sup> Not independent of ARTUSO 94  $B(h^- \pi^0 \nu_\tau)$  and BATTLE 94  $B(K^- \pi^0 \nu_\tau)$  values.
- <sup>3</sup> Not independent of BUSKULIC 96  $B(h^- \pi^0 \nu_\tau)$  and  $B(K^- \pi^0 \nu_\tau)$  values.
- <sup>4</sup> The authors divide by  $(\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10})/\Gamma = 0.467$  to obtain this result.
- <sup>5</sup> Experiment had no hadron identification. Kaon corrections were made, but insufficient information is given to permit their removal.
- <sup>6</sup> BURCHAT 87 value is not independent of YELTON 86 value. Nonresonant decays included.

 $\Gamma(\pi^- \pi^0 \text{ non-}\rho(770)\nu_\tau)/\Gamma_{total}$   $\Gamma_{15}/\Gamma$ 

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.3 ± 0.1 ± 0.3</b>	<sup>1</sup> BEHREND	84 CELL	$E_{cm}^{ee} = 14,22$ GeV

- <sup>1</sup> BEHREND 84 assume a flat nonresonant mass distribution down to the  $\rho(770)$  mass, using events with mass above 1300 to set the level.

 $\Gamma(K^- \pi^0 \nu_\tau)/\Gamma_{total}$   $\Gamma_{16}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.433 ± 0.015 OUR FIT</b>				
<b>0.426 ± 0.016 OUR AVERAGE</b>				
0.416 ± 0.003 ± 0.018	78k	AUBERT	07AP BABR	230 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV
0.471 ± 0.059 ± 0.023	360	ABBIENDI	04J OPAL	1991-1995 LEP runs
0.444 ± 0.026 ± 0.024	923	BARATE	99K ALEP	1991-1995 LEP runs
0.51 ± 0.10 ± 0.07	37	BATTLE	94 CLEO	$E_{cm}^{ee} \approx 10.6$ GeV
• • •				We do not use the following data for averages, fits, limits, etc. • • •
0.52 ± 0.04 ± 0.05	395	BUSKULIC	96 ALEP	Repl. by BARATE 99K





$$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{29}/\Gamma$$

$$\Gamma_{29}/\Gamma = (\Gamma_{30} + 0.3268\Gamma_{148} + 0.3268\Gamma_{152})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.16±0.04 OUR FIT**

**0.16±0.05±0.05**

<sup>1</sup> PROCARIO 93 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.16±0.04±0.09 232 <sup>2</sup> BUSKULIC 96 ALEP Repl. by SCHAEEL 05c

<sup>1</sup> PROCARIO 93 quotes  $B(h^- 4\pi^0 \nu_\tau)/B(h^- \pi^0 \nu_\tau) = 0.006 \pm 0.002 \pm 0.002$ . We multiply by the ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$  to obtain  $B(h^- 4\pi^0 \nu_\tau)$ . PROCARIO 93 assume  $B(h^- \geq 5 \pi^0 \nu_\tau)$  is small and do not correct for it.

<sup>2</sup> BUSKULIC 96 quote result for  $\tau^- \rightarrow h^- \geq 4\pi^0 \nu_\tau$ . We assume  $B(h^- \geq 5\pi^0 \nu_\tau)$  is negligible.

$$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}} \quad \Gamma_{30}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.11 ± 0.04 OUR FIT**

**0.112±0.037±0.035**

<sup>1</sup> SCHAEEL 05c ALEP 1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAEEL 05c  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{31}/\Gamma = (0.0021\Gamma_{20} + 0.0021\Gamma_{70})/\Gamma$$

The uncertainty on  $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau)/\Gamma_{\text{total}}$  takes into account the non-negligible contribution from the uncertainty of the coefficient of the relationship that defines  $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau)$  in terms of  $\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega))$ .

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**3.8±1.5 OUR FIT**

$$\Gamma(K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{32}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.552±0.029 OUR FIT**

**1.53 ± 0.04 OUR AVERAGE**

1.528±0.039±0.040 <sup>1</sup> ABBIENDI 01J OPAL 1990-1995 LEP runs

1.54 ± 0.24 ABREU 94K DLPH LEP 1992 Z data

1.70 ± 0.12 ± 0.19 202 <sup>2</sup> BATTLE 94 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

• • • We use the following data for averages but not for fits. • • •

1.520±0.040±0.041 4006 <sup>3</sup> BARATE 99K ALEP 1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.70 ± 0.05 ± 0.06 1610 <sup>4</sup> BUSKULIC 96 ALEP Repl. by BARATE 99K

1.6 ± 0.4 ± 0.2 35 AIHARA 87B TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

1.71 ± 0.29 53 MILLS 84 DLCO  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  is 0.60.

<sup>2</sup> BATTLE 94 quote  $1.60 \pm 0.12 \pm 0.19$ . We add  $0.10 \pm 0.02$  to correct for their rejection of  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>3</sup> Not independent of BARATE 99K  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ ,  $B(K^- 3\pi^0 \nu_\tau (\text{ex. } K^0))$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

<sup>4</sup> Not independent of BUSKULIC 96  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

$$\Gamma(K^- \geq 1(\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{33}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.859±0.028 OUR FIT**

**0.86 ± 0.05 OUR AVERAGE**

• • • We use the following data for averages but not for fits. • • •

0.869±0.031±0.034 <sup>1</sup> ABBIENDI 01J OPAL 1990-1995 LEP runs

0.69 ± 0.25 <sup>2</sup> ABREU 94K DLPH LEP 1992 Z data

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.2 ± 0.5 ± 0.2 <sup>9</sup> AIHARA 87B TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

<sup>1</sup> Not independent of ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  and  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$  values.

<sup>2</sup> Not independent of ABREU 94K  $B(K^- \nu_\tau)$  and  $B(K^- \geq 0 \text{ neutrals } \nu_\tau)$  measurements.

$$\Gamma(K_S^0 (\text{particles})^- \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{34}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.944±0.028 OUR FIT**

**0.918±0.015 OUR AVERAGE**

0.970±0.058±0.062 929 BARATE 98E ALEP 1991-1995 LEP runs

0.97 ± 0.09 ± 0.06 141 AKERS 94G OPAL  $E_{\text{cm}}^{\text{ee}} = 88-94$  GeV

• • • We use the following data for averages but not for fits. • • •

0.915±0.001±0.015 398k <sup>1</sup> RYU 14 BELL 669 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

<sup>1</sup> Not independent of RYU 14 measurements of  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \nu_\tau)$ , and  $B(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)$ .

$$\Gamma(h^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{35}/\Gamma = (\Gamma_{36} + \Gamma_{38})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.987±0.014 OUR FIT**

**0.90 ± 0.07 OUR AVERAGE**

0.855±0.036±0.073 1242 COAN 96 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

• • • We use the following data for averages but not for fits. • • •

1.01 ± 0.11 ± 0.07 555 <sup>1</sup> BARATE 98E ALEP 1991-1995 LEP runs

<sup>1</sup> Not independent of BARATE 98E  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$  and  $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$  values.

$$\Gamma(\pi^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{36}/\Gamma$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**8.40±0.14 OUR FIT**

**8.39±0.22 OUR AVERAGE** Error includes scale factor of 1.5. See the ideogram below.

8.32±0.02±0.16 158k <sup>1</sup> RYU 14 BELL 669 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

9.33±0.68±0.49 377 ABBIENDI 00C OPAL 1991-1995 LEP runs

9.28±0.45±0.34 937 <sup>2</sup> BARATE 99K ALEP 1991-1995 LEP runs

9.5 ± 1.5 ± 0.6 <sup>3</sup> ACCIARRI 95F L3 1991-1993 LEP runs

• • • We use the following data for averages but not for fits. • • •

8.55±1.17±0.66 509 <sup>4</sup> BARATE 98E ALEP 1991-1995 LEP runs

7.04±0.41±0.72 <sup>5</sup> COAN 96 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

8.08±0.04±0.26 53k EPIFANOV 07 BELL Repl. by RYU 14

7.9 ± 1.0 ± 0.9 98 <sup>6</sup> BUSKULIC 96 ALEP Repl. by BARATE 99K

<sup>1</sup> RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

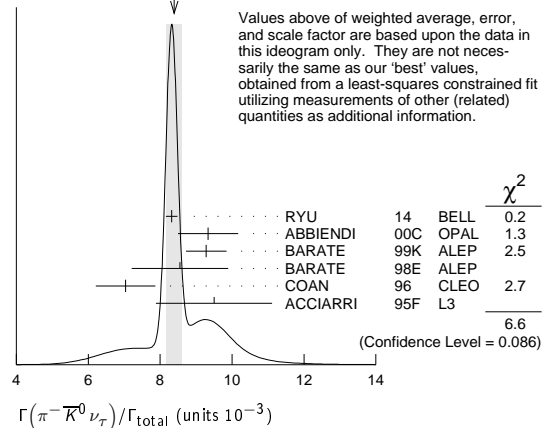
<sup>3</sup> ACCIARRI 95F do not identify  $\pi^-/K^-$  and assume  $B(K^- K^0 \nu_\tau) = (0.29 \pm 0.12)\%$ .

<sup>4</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. Not independent of BARATE 98E  $B(K^0 \text{ particles } \nu_\tau)$  value.

<sup>5</sup> Not independent of COAN 96  $B(h^- K^0 \nu_\tau)$  and  $B(K^- K^0 \nu_\tau)$  measurements.

<sup>6</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

WEIGHTED AVERAGE  
8.39±0.22 (Error scaled by 1.5)



$$\Gamma(\pi^- \bar{K}^0 (\text{non-} K^*(892)^- \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{37}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
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**5.4±2.1**

<sup>1</sup> EPIFANOV 07 BELL 351 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<17 95 ACCIARRI 95F L3 1991-1993 LEP runs

<sup>1</sup> EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) / B(\tau^- \rightarrow K_S^0 \pi^- \nu_\tau) = 0.933 \pm 0.027$ . We multiply their  $B(\tau^- \rightarrow \bar{K}^0 \pi^- \nu_\tau)$  by  $[1 - (0.933 \pm 0.027)]$  to obtain this result.

$$\Gamma(K^- K^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{38}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**14.8 ± 0.5 OUR FIT**

**14.9 ± 0.5 OUR AVERAGE**

14.80±0.14±0.54 33k <sup>1</sup> RYU 14 BELL 669 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

16.2 ± 2.1 ± 1.1 150 <sup>2</sup> BARATE 99K ALEP 1991-1995 LEP runs

15.8 ± 4.2 ± 1.7 46 <sup>3</sup> BARATE 98E ALEP 1991-1995 LEP runs

15.1 ± 2.1 ± 2.2 111 COAN 96 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

26 ± 9 ± 2 13 <sup>4</sup> BUSKULIC 96 ALEP Repl. by BARATE 99K

<sup>1</sup> RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>4</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

# Lepton Particle Listings

$\tau$

$\Gamma(K^-K^0\geq 0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{39}/\Gamma=(\Gamma_{38}+\Gamma_{43})/\Gamma$	
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.298±0.008 OUR FIT</b>					
<b>0.330±0.055±0.039</b>	124	ABBIENDI	00c	OPAL	1991–1995 LEP runs

$\Gamma(h^-\bar{K}^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$			$\Gamma_{40}/\Gamma=(\Gamma_{41}+\Gamma_{43})/\Gamma$		
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.532±0.013 OUR FIT</b>					
<b>0.50 ±0.06 OUR AVERAGE</b>		Error includes scale factor of 1.2.			
0.562±0.050±0.048	264	COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •					
0.446±0.052±0.046	157	<sup>1</sup> BARATE	98E	ALEP	1991–1995 LEP runs
<sup>1</sup> Not independent of BARATE 98E $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \tau)$ and $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$ values.					

$\Gamma(\pi^-K^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{41}/\Gamma$	
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.382±0.013 OUR FIT</b>					
<b>0.383±0.014 OUR AVERAGE</b>					
0.386±0.004±0.014	27k	<sup>1</sup> RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$
0.347±0.053±0.037	299	<sup>2</sup> BARATE	99k	ALEP	1991–1995 LEP runs
0.294±0.073±0.037	142	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
0.41 ±0.12 ±0.03		<sup>4</sup> ACCIARRI	95F	L3	1991–1993 LEP runs
• • • We use the following data for averages but not for fits. • • •					
0.417±0.058±0.044		<sup>5</sup> COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}}\approx 10.6\text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.32 ±0.11 ±0.05	23	<sup>6</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99k
<sup>1</sup> RYU 14 reconstruct $K^0$ 's using $K_S^0\rightarrow\pi^+\pi^-$ decays.					
<sup>2</sup> BARATE 99k measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					
<sup>3</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0\rightarrow\pi^+\pi^-$ decays.					
<sup>4</sup> ACCIARRI 95F do not identify $\pi^-/K^-$ and assume $B(K^-K^0\pi^0\nu_\tau)=(0.05\pm0.05)\%$ .					
<sup>5</sup> Not independent of COAN 96 $B(h^-K^0\pi^0\nu_\tau)$ and $B(K^-K^0\pi^0\nu_\tau)$ measurements.					
<sup>6</sup> BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					

$\Gamma(K^0\rho^-\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{42}/\Gamma$	
VALUE (%)		DOCUMENT ID	TECN	COMMENT	
<b>0.22 ±0.05 OUR AVERAGE</b>					
0.250±0.057±0.044		<sup>1</sup> BARATE	99k	ALEP	1991–1995 LEP runs
0.188±0.054±0.038		<sup>2</sup> BARATE	98E	ALEP	1991–1995 LEP runs
<sup>1</sup> BARATE 99k measure $K^0$ 's by detecting $K_L^0$ 's in hadron calorimeter. They determine the $K^0\rho^-$ fraction in $\tau^-\rightarrow\pi^-K^0\pi^0\nu_\tau$ decays to be $(0.72\pm0.12\pm0.10)$ and multiply their $B(\pi^-K^0\pi^0\nu_\tau)$ measurement by this fraction to obtain the quoted result.					
<sup>2</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0\rightarrow\pi^+\pi^-$ decays. They determine the $K^0\rho^-$ fraction in $\tau^-\rightarrow\pi^-K^0\pi^0\nu_\tau$ decays to be $(0.64\pm0.09\pm0.10)$ and multiply their $B(\pi^-K^0\pi^0\nu_\tau)$ measurement by this fraction to obtain the quoted result.					

$\Gamma(K^-K^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{43}/\Gamma$	
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>15.0 ±0.7 OUR FIT</b>					
<b>14.9 ±0.7 OUR AVERAGE</b>					
14.96±0.20±0.74	8.3k	<sup>1</sup> RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$
14.3 ±2.5 ±1.5	78	<sup>2</sup> BARATE	99k	ALEP	1991–1995 LEP runs
15.2 ±7.6 ±2.1	15	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
14.5 ±3.6 ±2.0	32	COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}}\approx 10.6\text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
10 ±5 ±3	5	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99k
<sup>1</sup> RYU 14 reconstruct $K^0$ 's using $K_S^0\rightarrow\pi^+\pi^-$ decays.					
<sup>2</sup> BARATE 99k measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					
<sup>3</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0\rightarrow\pi^+\pi^-$ decays.					
<sup>4</sup> BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					

$\Gamma(\pi^-K^0\geq 1\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{44}/\Gamma=(\Gamma_{41}+\Gamma_{45})/\Gamma$	
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.408±0.025 OUR FIT</b>					
<b>0.324±0.074±0.066</b>	148	ABBIENDI	00c	OPAL	1991–1995 LEP runs

$\Gamma(\pi^-K^0\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$				$\Gamma_{45}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.26±0.23 OUR FIT</b>					
<b>0.26±0.24</b>			<sup>1</sup> BARATE	99R	ALEP 1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.66	95	17	<sup>2</sup> BARATE	99k	ALEP 1991–1995 LEP runs
0.58±0.33±0.14		5	<sup>3</sup> BARATE	98E	ALEP 1991–1995 LEP runs
<sup>1</sup> BARATE 99R combine the BARATE 98E and BARATE 99k measurements to obtain this value.					
<sup>2</sup> BARATE 99k measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					
<sup>3</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0\rightarrow\pi^+\pi^-$ decays.					

$\Gamma(K^-K^0\pi^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{46}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.16 × 10<sup>-3</sup></b>	95	<sup>1</sup> BARATE	99R	ALEP	1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.18 × 10 <sup>-3</sup>	95	<sup>2</sup> BARATE	99k	ALEP	1991–1995 LEP runs
<0.39 × 10 <sup>-3</sup>	95	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
<sup>1</sup> BARATE 99R combine the BARATE 98E and BARATE 99k bounds to obtain this value.					
<sup>2</sup> BARATE 99k measure $K^0$ 's by detecting $K_L^0$ 's in hadron calorimeter.					
<sup>3</sup> BARATE 98E reconstruct $K^0$ 's by using $K_S^0\rightarrow\pi^+\pi^-$ decays.					

$\Gamma(\pi^-K^0\bar{K}^0\nu_\tau)/\Gamma_{\text{total}}$			$\Gamma_{47}/\Gamma=(\Gamma_{48}+\Gamma_{49}+\Gamma_{50})/\Gamma$		
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.155±0.024 OUR FIT</b>					
• • • We use the following data for averages but not for fits. • • •					
<b>0.153±0.030±0.016</b>	74	<sup>1</sup> BARATE	98E	ALEP	1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.31 ±0.12 ±0.04		<sup>2</sup> ACCIARRI	95F	L3	1991–1993 LEP runs
<sup>1</sup> BARATE 98E obtain this value by adding twice their $B(\pi^-K_S^0K_S^0\nu_\tau)$ value to their $B(\pi^-K_S^0K_L^0\nu_\tau)$ value.					
<sup>2</sup> ACCIARRI 95F assume $B(\pi^-K_S^0K_S^0\nu)=B(\pi^-K_S^0K_L^0\nu)=1/2B(\pi^-K_S^0K_L^0\nu)$ .					

$\Gamma(\pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{48}/\Gamma$	
Bose-Einstein correlations might make the mixing fraction different than 1/4.					
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.33±0.07 OUR FIT</b>					
<b>2.32±0.06 OUR AVERAGE</b>					
2.33±0.03±0.09	6.7k	RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}$ =10.6 GeV
2.31±0.04±0.08	5.0k	LEES	12Y	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}$ =10.6 GeV
2.6 ±1.0 ±0.5	6	BARATE	98E	ALEP	1991–1995 LEP runs
2.3 ±0.5 ±0.3	42	COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}}$ ≈ 10.6 GeV

$\Gamma(\pi^-K_S^0K_L^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{49}/\Gamma$	
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>10.8±2.4 OUR FIT</b>					
<b>10.1±2.3±1.3</b>	68	BARATE	98E	ALEP	1991–1995 LEP runs
$\Gamma(\pi^-K_L^0K_L^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{50}/\Gamma=\Gamma_{48}/\Gamma$	
VALUE (units 10 <sup>-4</sup> )		DOCUMENT ID			
<b>2.33±0.07 OUR FIT</b>					

$\Gamma(\pi^-K^0\bar{K}^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{51}/\Gamma=(\Gamma_{52}+\Gamma_{56}+\Gamma_{57})/\Gamma$		
VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID	TECN	COMMENT
<b>3.6±1.2 OUR FIT</b>			
• • • We use the following data for averages but not for fits. • • •			
<b>3.1±2.3</b>	<sup>1</sup> BARATE	99R	ALEP 1991-1995 LEP runs
<sup>1</sup> BARATE 99R combine	BARATE 98E	$\Gamma(\pi^-K_S^0K_S^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$	and
$\Gamma(\pi^-K_S^0K_L^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$	measurements to obtain this value.		

$\Gamma(\pi^-K_S^0K_S^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{52}/\Gamma$	
VALUE (units 10 <sup>-5</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.82±0.21 OUR FIT</b>					
<b>1.80±0.21 OUR AVERAGE</b>					
2.00±0.22±0.20	303	RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$
1.60±0.20±0.22	409	LEES	12Y	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<20	95	BARATE	98E	ALEP	1991–1995 LEP runs

$\Gamma(K^{*-}K^0\pi^0\nu_\tau\rightarrow\pi^-K_S^0K_S^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{53}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )		DOCUMENT ID	TECN	COMMENT	
<b>10.8±1.4±1.5</b>		RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$

$\Gamma(f_1(1285)\pi^-\nu_\tau\rightarrow\pi^-K_S^0K_S^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{54}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )		DOCUMENT ID	TECN	COMMENT	
<b>6.8±1.3±0.7</b>		RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$

$\Gamma(f_1(1420)\pi^-\nu_\tau\rightarrow\pi^-K_S^0K_S^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{55}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )		DOCUMENT ID	TECN	COMMENT	
<b>2.4±0.5±0.6</b>		RYU	14	BELL	669 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6\text{ GeV}$

$\Gamma(\pi^-K_S^0K_L^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{56}/\Gamma$	
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3.2±1.2 OUR FIT</b>					
<b>3.1±1.1±0.5</b>	11	BARATE	98E	ALEP	1991–1995 LEP runs

$\Gamma(\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{57}/\Gamma = \Gamma_{52}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID
<b>1.82 ± 0.21 OUR FIT</b>	

$\Gamma(K^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{58}/\Gamma$
VALUE	CL%
<b>&lt;6.3 × 10<sup>-7</sup></b>	90
DOCUMENT ID	TECN
LEES	12Y BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

$\Gamma(K^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{59}/\Gamma$
VALUE	CL%
<b>&lt;4.0 × 10<sup>-7</sup></b>	90
DOCUMENT ID	TECN
LEES	12Y BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

$\Gamma(K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{60}/\Gamma$
VALUE (%)	CL%
<b>&lt;0.17</b>	95
DOCUMENT ID	TECN
TSCIRHART	88 HRS $E_{\text{cm}}^{\text{ee}} = 29$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<0.27	90
BELTRAMI	85 HRS $E_{\text{cm}}^{\text{ee}} = 29$ GeV

$\Gamma(K^0 h^+ h^- h^- \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{61}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b>2.5 ± 2.0 OUR FIT</b>	
<b>2.3 ± 1.9 ± 0.7</b>	6
DOCUMENT ID	TECN
1 BARATE	98E ALEP 1991–1995 LEP runs
1 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.	

$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{62}/\Gamma$
$\Gamma_{62}/\Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.4247\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.2628\Gamma_{154} + 0.7259\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178})/\Gamma$	

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>15.21 ± 0.06 OUR FIT</b>				
<b>14.8 ± 0.4 OUR AVERAGE</b>				
14.4 ± 0.6 ± 0.3		ADEVA	91F L3	$E_{\text{cm}}^{\text{ee}} = 88.3\text{--}94.3$ GeV
15.0 ± 0.4 ± 0.3		BEHREND	89B CELL	$E_{\text{cm}}^{\text{ee}} = 14\text{--}47$ GeV
15.1 ± 0.8 ± 0.6		AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13.5 ± 0.3 ± 0.3		ABACHI	89B HRS	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
12.8 ± 1.0 ± 0.7		1 BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
12.1 ± 0.5 ± 1.2		RUCKSTUHL	86 DLCO	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
12.8 ± 0.5 ± 0.8	1420	SCHMIDKE	86 MRK2	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
15.3 ± 1.1 $\pm_{-1.6}^{+1.3}$	367	ALTHOFF	85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5$ GeV
13.6 ± 0.5 ± 0.8		BARTEL	85F JADE	$E_{\text{cm}}^{\text{ee}} = 34.6$ GeV
12.2 ± 1.3 ± 3.9		2 BERGER	85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6$ GeV
13.3 ± 0.3 ± 0.6		FERNANDEZ	85 MAC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
24 ± 6	35	BRANDELIK	80 TASS	$E_{\text{cm}}^{\text{ee}} = 30$ GeV
32 ± 5	692	3 BACINO	78B DLCO	$E_{\text{cm}}^{\text{ee}} = 3.1\text{--}7.4$ GeV
35 ± 11		3 BRANDELIK	78 DASP	Assumes $V\text{--}A$ decay
18 ± 6.5	33	3 JAROS	78 LGW	$E_{\text{cm}}^{\text{ee}} > 6$ GeV

1 BURCHAT 87 value is not independent of SCHMIDKE 86 value.

2 Not independent of BERGER 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ , and therefore not used in the fit.

3 Low energy experiments are not in average or fit because the systematic errors in background subtraction are judged to be large.

$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-) (\text{"3-prong"})/\Gamma_{\text{total}}$	$\Gamma_{63}/\Gamma$
$\Gamma_{63}/\Gamma = (\Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.489\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178})/\Gamma$	

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.55 ± 0.06 OUR FIT</b>				
<b>14.61 ± 0.06 OUR AVERAGE</b>				
14.556 ± 0.105 ± 0.076		1 ACHARD	01D L3	1992–1995 LEP runs
14.96 ± 0.09 ± 0.22	10.4k	AKERS	95Y OPAL	1991–1994 LEP runs
• • • We use the following data for averages but not for fits. • • •				
14.652 ± 0.067 ± 0.086		SCHAE	05C ALEP	1991–1995 LEP runs
14.569 ± 0.093 ± 0.048	23k	2 ABREU	01M DLPH	1992–1995 LEP runs
14.22 ± 0.10 ± 0.37		3 BALEST	95C CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
15.26 ± 0.26 ± 0.22		ACTON	92H OPAL	Repl. by AKERS 95Y
13.3 ± 0.3 ± 0.8		4 ALBRECHT	92D ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV
14.35 $\pm_{-0.45}^{+0.40}$ ± 0.24		DECAMP	92C ALEP	1989–1990 LEP runs

1 The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.19$  respectively.

2 The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{1-prong})$  and  $B(\tau \rightarrow \text{5-prong})$  are  $-0.98$  and  $-0.08$  respectively.

3 Not independent of BALEST 95C  $B(h^- h^- h^+ \nu_\tau)$  and  $B(h^- h^- h^+ \pi^0 \nu_\tau)$  values, and BORTOLETTO 93  $B(h^- h^- h^+ 2\pi^0 \nu_\tau)/B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau)$  value.

4 This ALBRECHT 92D value is not independent of their  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  value.

$\Gamma(h^- h^- h^+ \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{64}/\Gamma$
$\Gamma_{64}/\Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + \Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489\Gamma_{168} + 0.0153\Gamma_{176} + 0.0153\Gamma_{177})/\Gamma$	

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.80 ± 0.05 OUR FIT</b>				
• • • We use the following data for averages but not for fits. • • •				
<b>7.6 ± 0.1 ± 0.5</b>	7.5k	1 ALBRECHT	96E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
9.92 ± 0.10 ± 0.09	11.2k	2 BUSKULIC	96 ALEP	Repl. by SCHAE 05C
9.49 ± 0.36 ± 0.63		DECAMP	92C ALEP	Repl. by SCHAE 05C
8.7 ± 0.7 ± 0.3	694	3 BEHREND	90 CELL	$E_{\text{cm}}^{\text{ee}} = 35$ GeV
7.0 ± 0.3 ± 0.7	1566	4 BAND	87 MAC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
6.7 ± 0.8 ± 0.9		5 BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
6.4 ± 0.4 ± 0.9		6 RUCKSTUHL	86 DLCO	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
7.8 ± 0.5 ± 0.8	890	SCHMIDKE	86 MRK2	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
8.4 ± 0.4 ± 0.7	1255	6 FERNANDEZ	85 MAC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
9.7 ± 2.0 ± 1.3		BEHREND	84 CELL	$E_{\text{cm}}^{\text{ee}} = 14.22$ GeV

1 ALBRECHT 96E not independent of ALBRECHT 93C  $\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0) \times \Gamma(\text{particle}^- \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$  value.

2 BUSKULIC 96 quote  $B(h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) = 9.50 \pm 0.10 \pm 0.11$ . We add 0.42 to remove their  $K^0$  correction and reduce the systematic error accordingly.

3 BEHREND 90 subtract 0.3% to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution to measured events.

4 BAND 87 subtract for charged kaon modes; not independent of FERNANDEZ 85 value.

5 BURCHAT 87 value is not independent of SCHMIDKE 86 value.

6 Value obtained by multiplying paper's  $R = B(h^- h^- h^+ \nu_\tau)/B(3\text{-prong})$  by  $B(3\text{-prong}) = 0.143$  and subtracting 0.3% for  $K^*(892)$  background.

$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$	$\Gamma_{65}/\Gamma$
$\Gamma_{65}/\Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489\Gamma_{168} + 0.0153\Gamma_{176} + 0.0153\Gamma_{177})/\Gamma$	

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.46 ± 0.05 OUR FIT</b>				
<b>9.44 ± 0.14 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
9.317 ± 0.090 ± 0.082	12.2k	1 ABDALLAH	06A DLPH	1992–1995 LEP runs
9.51 ± 0.07 ± 0.20	37.7k	BALEST	95C CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV

• • • We use the following data for averages but not for fits. • • •

9.87 ± 0.10 ± 0.24

2 AKERS 95Y OPAL 1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.50 ± 0.10 ± 0.11

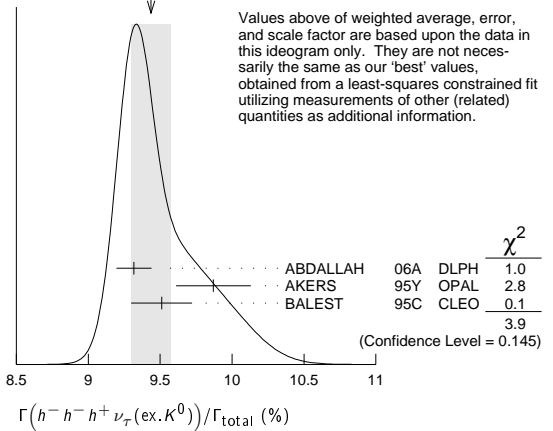
3 BUSKULIC 96 ALEP Repl. by SCHAE 05C

1 See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

2 Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

3 Not independent of BUSKULIC 96  $B(h^- h^- h^+ \nu_\tau)$  value.

WEIGHTED AVERAGE  
9.44 ± 0.14 (Error scaled by 1.4)



$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-) (\text{"3-prong"}))$	$\Gamma_{65}/\Gamma_{63}$
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$$\Gamma_{65}/\Gamma_{63} = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489\Gamma_{168} + 0.0153\Gamma_{176} + 0.0153\Gamma_{177}) / (0.4247\Gamma_{52} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2292\Gamma_{149} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.1131\Gamma_{154} + 0.3268\Gamma_{158} + 0.489\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178} + 0.892\Gamma_{180})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>64.98 ± 0.31 OUR FIT</b>			
<b>66.0 ± 0.4 ± 1.4</b>	AKERS	95Y OPAL	1991–1994 LEP runs

# Lepton Particle Listings

$\tau$

$$\Gamma(h^-h^-h^+\nu_\tau(\text{ex.}K^0\omega))/\Gamma_{\text{total}} \quad \Gamma_{66}/\Gamma$$
$$\Gamma_{66}/\Gamma = (\Gamma_{70}+\Gamma_{97}+\Gamma_{106}+0.489\Gamma_{168})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>9.43±0.05 OUR FIT</b>	

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{67}/\Gamma = (0.34598\Gamma_{36}+\Gamma_{70}+0.0153\Gamma_{176})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>9.31±0.05 OUR FIT</b>	

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{68}/\Gamma = (\Gamma_{70}+0.0153\Gamma_{176})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.02±0.05 OUR FIT</b>				
<b>8.77±0.13 OUR AVERAGE</b>		Error includes scale factor of 1.1.		

8.42±0.00 <sub>-0.25</sub> <sup>+0.26</sup>	8.9M	1 LEE	10	BELL	666 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
8.83±0.01±0.13	1.6M	2 AUBERT	08	BABR	342 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
9.13±0.05±0.46	43k	3 BRIERE	03	CLE3		$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> Quoted statistical error is 0.003%. Correlation matrix for LEE 10 branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$

(1)	(2)	(3)
(2)	0.175	
(3)	0.049	0.080
(4)	-0.053	0.035 -0.008

<sup>2</sup> Correlation matrix for AUBERT 08 branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$

(1)	(2)	(3)
(2)	0.544	
(3)	0.390	0.177
(4)	0.031	0.093 0.087

<sup>3</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  and 71% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0, \text{non-axial vector}))/\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)) \quad \Gamma_{69}/\Gamma_{68}$$
$$\Gamma_{69}/\Gamma_{68} = \Gamma_{69}/(\Gamma_{70}+0.0153\Gamma_{175})$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.261</b>	95	1 ACKERSTAFF	97R	OPAL 1992-1994 LEP runs

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $\text{B}(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.}K^0))$  from non-axial vectors.

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}} \quad \Gamma_{70}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.99 ±0.05 OUR FIT</b>				
<b>9.041±0.060±0.076</b>	29k	1 SCHAE	05c	ALEP 1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAE 05c  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^-h^-h^+ \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{71}/\Gamma$$
$$\Gamma_{71}/\Gamma = (0.34598\Gamma_{41}+0.34598\Gamma_{43}+0.4247\Gamma_{48}+0.4247\Gamma_{52}+\Gamma_{78}+\Gamma_{85}+\Gamma_{86}+0.103+\Gamma_{107}+0.2810\Gamma_{148}+0.2810\Gamma_{150}+0.2810\Gamma_{152}+0.2926\Gamma_{154}+0.892\Gamma_{176}+0.892\Gamma_{177}+0.9078\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.29±0.05 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

5.6 ±0.7 ±0.3	352	1 BEHREND	90	CELL		$E_{\text{cm}}^{\text{ee}} = 35$ GeV
4.2 ±0.5 ±0.9	203	2 ALBRECHT	87L	ARG		$E_{\text{cm}}^{\text{ee}} = 10$ GeV
6.1 ±0.8 ±0.9		3 BURCHAT	87	MRK2		$E_{\text{cm}}^{\text{ee}} = 29$ GeV
7.6 ±0.4 ±0.9		4,5 RUCKSTUHL	86	DLEO		$E_{\text{cm}}^{\text{ee}} = 29$ GeV
4.7 ±0.5 ±0.8	530	6 SCHMIDKE	86	MRK2		$E_{\text{cm}}^{\text{ee}} = 29$ GeV
5.6 ±0.4 ±0.7		5 FERNANDEZ	85	MAC		$E_{\text{cm}}^{\text{ee}} = 29$ GeV
6.2 ±2.3 ±1.7		BEHREND	84	CELL		$E_{\text{cm}}^{\text{ee}} = 14,22$ GeV

<sup>1</sup> BEHREND 90 value is not independent of BEHREND 90  $\text{B}(3h\nu_\tau \geq 1 \text{ neutrals}) + \text{B}(5\text{-prong})$ .

<sup>2</sup> ALBRECHT 87L measure the product of branching ratios  $\text{B}(3\pi^\pm \pi^0 \nu_\tau) \text{B}((e\bar{\nu}_e \text{ or } \mu\bar{\nu}_e \text{ or } \pi^0 \text{ or } \rho) \nu_\tau) = 0.029$  and use the PDG 86 values for the second branching ratio which sum to 0.69 ± 0.03 to get the quoted value.

<sup>3</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>4</sup> Contributions from kaons and from  $>1\pi^0$  are subtracted. Not independent of (3-prong +  $0\pi^0$ ) and (3-prong +  $\geq 0\pi^0$ ) values.

<sup>5</sup> Value obtained using paper's  $R = \text{B}(h^-h^-h^+\nu_\tau)/\text{B}(3\text{-prong})$  and current  $\text{B}(3\text{-prong}) = 0.143$ .

<sup>6</sup> Not independent of SCHMIDKE 86  $h^-h^-h^+\nu_\tau$  and  $h^-h^-h^+(\geq 0\pi^0)\nu_\tau$  values.

$$\Gamma(h^-h^-h^+ \geq 1\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{72}/\Gamma$$
$$\Gamma_{72}/\Gamma = (\Gamma_{78}+\Gamma_{85}+\Gamma_{86}+\Gamma_{103}+\Gamma_{107}+0.2292\Gamma_{148}+0.2292\Gamma_{150}+0.2292\Gamma_{152}+0.892\Gamma_{176}+0.892\Gamma_{177}+0.9078\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.09 ±0.05 OUR FIT</b>				
<b>5.10 ±0.12 OUR AVERAGE</b>				

• • • We use the following data for averages but not for fits. • • •

5.106±0.083±0.103	10.1k	1 ABDALLAH	06A	DLPH	1992-1995 LEP runs
5.09 ±0.10 ±0.23		2 AKERS	95Y	OPAL	1991-1994 LEP runs
4.95 ±0.29 ±0.65	570	DECAMP	92c	ALEP	Repl. by SCHAE 05c

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of AKERS 95Y  $\text{B}(h^-h^-h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex.}K_S^0 \rightarrow \pi^+\pi^-))$  and  $\text{B}(h^-h^-h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex.}K^0))/\text{B}(h^-h^-h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex.}K_S^0 \rightarrow \pi^+\pi^-))$  values.

$$\Gamma(h^-h^-h^+\pi^0\nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{73}/\Gamma$$
$$\Gamma_{73}/\Gamma = (0.34598\Gamma_{41}+0.34598\Gamma_{43}+\Gamma_{78}+\Gamma_{103}+\Gamma_{107}+0.2292\Gamma_{150}+0.892\Gamma_{176}+0.892\Gamma_{177}+0.0153\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.76±0.05 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.45±0.09±0.07	6.1k	1 BUSKULIC	96	ALEP	Repl. by SCHAE 05c
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<sup>1</sup> BUSKULIC 96 quote  $\text{B}(h^-h^-h^+\pi^0\nu_\tau(\text{ex.}K^0)) = 4.30 \pm 0.09 \pm 0.09$ . We add 0.15 to remove their  $K^0$  correction and reduce the systematic error accordingly.

$$\Gamma(h^-h^-h^+\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{74}/\Gamma$$
$$\Gamma_{74}/\Gamma = (\Gamma_{78}+\Gamma_{103}+\Gamma_{107}+0.2292\Gamma_{150}+0.892\Gamma_{176}+0.892\Gamma_{177}+0.0153\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.57 ±0.05 OUR FIT</b>				
<b>4.45 ±0.14 OUR AVERAGE</b>		Error includes scale factor of 1.2.		

4.545±0.106±0.103	8.9k	1 ABDALLAH	06A	DLPH	1992-1995 LEP runs
4.23 ±0.06 ±0.22	7.2k	BALEST	95c	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^-h^-h^+\pi^0\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}} \quad \Gamma_{75}/\Gamma = (\Gamma_{78}+\Gamma_{103}+\Gamma_{107}+0.2292\Gamma_{150})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>2.79±0.07 OUR FIT</b>	

$$\Gamma(\pi^-\pi^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{76}/\Gamma$$
$$\Gamma_{76}/\Gamma = (0.34598\Gamma_{41}+\Gamma_{78}+0.892\Gamma_{176}+0.0153\Gamma_{178})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>4.62±0.05 OUR FIT</b>	

$$\Gamma(\pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{77}/\Gamma$$
$$\Gamma_{77}/\Gamma = (\Gamma_{78}+0.892\Gamma_{176}+0.0153\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.49 ±0.05 OUR FIT</b>				
<b>4.55 ±0.13 OUR AVERAGE</b>		Error includes scale factor of 1.6.		

4.598±0.057±0.064	16k	1 SCHAE	05c	ALEP	1991-1995 LEP runs
4.19 ±0.10 ±0.21		2 EDWARDS	00A	CLEO	4.7 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> SCHAE 05c quote  $(4.590 \pm 0.057 \pm 0.064)\%$ . We add 0.008% to remove their correction for  $\tau^- \rightarrow \pi^- \pi^0 \omega \nu_\tau \rightarrow \pi^- \pi^0 \pi^+ \pi^- \nu_\tau$  decays. See footnote to SCHAE 05c

$\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> EDWARDS 00A quote  $(4.19 \pm 0.10) \times 10^{-2}$  with a 5% systematic error.

$$\Gamma(\pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}} \quad \Gamma_{78}/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>2.74±0.07 OUR FIT</b>	

$$\Gamma(h^- \rho \pi^0 \nu_\tau)/\Gamma(h^-h^-h^+\pi^0\nu_\tau) \quad \Gamma_{79}/\Gamma_{73}$$

• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.30±0.04±0.02	393	ALBRECHT	91D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV

$$\Gamma(h^- \rho^+ h^- \nu_\tau)/\Gamma(h^-h^-h^+\pi^0\nu_\tau) \quad \Gamma_{80}/\Gamma_{73}$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$$\Gamma(h^- \rho^- h^+ \nu_\tau)/\Gamma(h^-h^-h^+\pi^0\nu_\tau) \quad \Gamma_{81}/\Gamma_{73}$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

• • • We do not use the following data for averages, fits, limits, etc. • • •

See key on page 601

## Lepton Particle Listings

T

$$\Gamma(h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{82}/\Gamma$$

$$\Gamma_{82}/\Gamma = (\Gamma_{85} + \Gamma_{86} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.517±0.031 OUR FIT**

1.3k 1 ABDALLAH 06A DLPH 1992-1995 LEP runs

1 See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{83}/\Gamma$$

$$\Gamma_{83}/\Gamma = (0.4247\Gamma_{48} + \Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/\Gamma$$

VALUE (%)	DOCUMENT ID
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**0.505±0.031 OUR FIT**

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{84}/\Gamma$$

$$\Gamma_{84}/\Gamma = (\Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.495±0.031 OUR FIT**

2.6k 1 SCHAEEL 05c ALEP 1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.50 ± 0.07 ± 0.07 1.8k BUSKULIC 96 ALEP Repl. by SCHAEEL 05c

1 SCHAEEL 05c quote (0.392 ± 0.030 ± 0.035)%. We add 0.043% to remove their correction for  $\tau^- \rightarrow \pi^- \eta \pi^0 \nu_\tau \rightarrow \pi^- \pi^+ \pi^- 2\pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow K^- \pi^+ \pi^- 2\pi^0 \nu_\tau$  decays. See footnote to SCHAEEL 05c  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) \quad \Gamma_{84}/\Gamma_{62}$$

$$\Gamma_{84}/\Gamma_{62} = (\Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/(0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2292\Gamma_{149} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.3759\Gamma_{154} + 0.3268\Gamma_{158} + 0.7259\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178} + 0.892\Gamma_{180})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**3.26±0.20 OUR FIT**3.4 ± 0.2 ± 0.3 668 BORTOLETTO93 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0 \omega, \eta))/\Gamma_{\text{total}} \quad \Gamma_{85}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
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**10±4 OUR FIT**

$$\Gamma(h^- h^- h^+ 3\pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{86}/\Gamma = (0.4247\Gamma_{52} + \Gamma_{87} + 0.1131\Gamma_{154})/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**2.12±0.30 OUR FIT**2.2 ± 0.3 ± 0.4 139 ANASTASSOV 01 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

&lt; 4.9 95 SCHAEEL 05c ALEP 1991-1995 LEP runs

2.85 ± 0.56 ± 0.51 57 ANDERSON 97 CLEO Repl. by ANASTASSOV 01

11 ± 4 ± 5 440 1 BUSKULIC 96 ALEP Repl. by SCHAEEL 05c

1 BUSKULIC 96 state their measurement is for  $B(h^- h^- h^+ \geq 3\pi^0 \nu_\tau)$ . We assume that  $B(h^- h^- h^+ \geq 4\pi^0 \nu_\tau)$  is very small.

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{87}/\Gamma$$

$$\Gamma_{87}/\Gamma = (\Gamma_{89} + 0.2292\Gamma_{149} + 0.3268\Gamma_{158} + 0.892\Gamma_{180})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
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**1.94±0.30 OUR FIT**

• • • We use the following data for averages but not for fits. • • •

2.07±0.18±0.37 1 LEES 12x BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV1 Not independent of LEES 12x  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^0 \pi^0 \nu_\tau)/\Gamma$ ,  $\Gamma(\tau^- \rightarrow \pi^- \omega 2\pi^0 \nu_\tau)/\Gamma$ , and  $\Gamma(\tau^- \rightarrow \bar{\eta}_1(1285) \pi^- \nu_\tau \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau)/\Gamma$  values.

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \bar{\eta}_1(1285)))/\Gamma_{\text{total}} \quad \Gamma_{88}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
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**1.69±0.08±0.43**LEES 12x BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, \bar{\eta}_1(1285)))/\Gamma_{\text{total}} \quad \Gamma_{89}/\Gamma$$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
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**1.4±2.7 OUR FIT**1.0 ± 0.8 ± 3.0 1 LEES 12x BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV1 LEES 12x measurement corresponds to the lower limit of  $< 5.8 \times 10^{-5}$  at 90% CL.

$$\Gamma(K^- h^+ h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{90}/\Gamma$$

$$\Gamma_{90}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{150} + 0.489\Gamma_{168} + 0.9078\Gamma_{177})/\Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
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**0.629±0.014 OUR FIT**<0.6 90 AIHARA 84c TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{91}/\Gamma = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{177})/\Gamma$$

VALUE (%)	DOCUMENT ID
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**0.437±0.007 OUR FIT**

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{91}/\Gamma_{68}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**4.84±0.08 OUR FIT**5.44 ± 0.21 ± 0.53 7.9k RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{92}/\Gamma$$

$$\Gamma_{92}/\Gamma = (\Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
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**8.6±1.2 OUR FIT**

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{92}/\Gamma_{77}$$

$$\Gamma_{92}/\Gamma_{77} = (\Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/(\Gamma_{78} + 0.892\Gamma_{176} + 0.0153\Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.91±0.26 OUR FIT**2.61 ± 0.45 ± 0.42 719 RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{93}/\Gamma$$

$$\Gamma_{93}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + 0.2810\Gamma_{150} + 0.9078\Gamma_{177})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.477±0.014 OUR FIT**0.58 ± 0.15 ± 0.12 20 1 BAUER 94 TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.22 ± 0.16 ± 0.05 9 2 MILLS 85 DLCO  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

1 We multiply 0.58% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

2 Error correlated with MILLS 85 ( $K K \pi \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{94}/\Gamma$$

$$\Gamma_{94}/\Gamma = (\Gamma_{97} + \Gamma_{103} + 0.2292\Gamma_{150} + 0.9078\Gamma_{177})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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**0.373±0.013 OUR FIT****0.30 ± 0.05 OUR AVERAGE**

• • • We use the following data for averages but not for fits. • • •

0.343 ± 0.073 ± 0.031 ABBIENDI 00D OPAL 1990-1995 LEP runs

0.275 ± 0.064 1 BARATE 98 ALEP 1991-1995 LEP runs

1 Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{95}/\Gamma = (0.34598\Gamma_{38} + \Gamma_{97} + 0.0153\Gamma_{177})/\Gamma$$

VALUE (%)	DOCUMENT ID
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**0.345±0.007 OUR FIT**

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{96}/\Gamma = (\Gamma_{97} + 0.0153\Gamma_{177})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.293±0.007 OUR FIT****0.290±0.018 OUR AVERAGE**

Error includes scale factor of 2.4. See the ideogram below.

0.330 ± 0.001 ± 0.016 ± 0.017 794k 1 LEE 10 BELL 666 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV0.273 ± 0.002 ± 0.009 70k 2 AUBERT 08 BABR 342 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

0.415 ± 0.053 ± 0.040 269 ABBIENDI 04J OPAL 1991-1995 LEP runs

0.384 ± 0.014 ± 0.038 3.5k 3 BRIERE 03 CLE3  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

0.214 ± 0.037 ± 0.029 BARATE 98 ALEP 1991-1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

0.346 ± 0.023 ± 0.056 158 4 RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

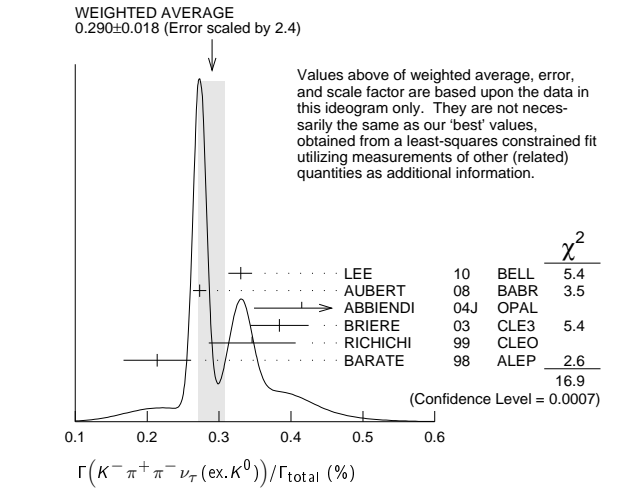
0.360 ± 0.082 ± 0.048 ABBIENDI 00D OPAL 1990-1995 LEP runs

1 See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  value.2 See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements.3 47% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

4 Not independent of RICHICHI 99

 $\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ ,  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95c  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.

## Lepton Particle Listings

 $\tau$ 

$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)) \quad \Gamma_{96}/\Gamma_{68}$$

$$\Gamma_{96}/\Gamma_{68} = (\Gamma_{97} + 0.0153\Gamma_{177})/(\Gamma_{70} + 0.0153\Gamma_{176})$$

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.25±0.07 OUR FIT</b>				
• • • We use the following data for averages but not for fits. • • •				

3.92±0.02±0.15 0.16	794k	<sup>1</sup> LEE	10	BELL	666 fb <sup>-1</sup>	E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
<sup>1</sup> Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ values.						

$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}} \quad \Gamma_{97}/\Gamma$$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID
<b>2.93±0.07 OUR FIT</b>	

$$\Gamma(K^-\rho^0\nu_\tau \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)) \quad \Gamma_{98}/\Gamma_{96}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.48±0.14±0.10</b>	<sup>1</sup> ASNER	00B	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.39±0.14	<sup>2</sup> BARATE	99R	ALEP 1991-1995 LEP runs

<sup>1</sup> ASNER 00B assume  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)$  decays proceed only through  $K\rho$  and  $K^*\pi$  intermediate states. They assume the resonance structure of  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)$  decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances, and assume  $B(K_1(1270) \rightarrow K^*(892)\pi) = (16 \pm 5)\%$ ,  $B(K_1(1270) \rightarrow K\rho) = (42 \pm 6)\%$ , and  $B(K_1(1400) \rightarrow K\rho) = 0$ .

<sup>2</sup> BARATE 99R assume  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)$  decays proceed only through  $K\rho$  and  $K^*\pi$  intermediate states. The quoted error is statistical only.

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{99}/\Gamma$$

$$\Gamma_{99}/\Gamma = (0.3459\Gamma_{43} + \Gamma_{103} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID
<b>13.1±1.2 OUR FIT</b>	

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{100}/\Gamma$$

$$\Gamma_{100}/\Gamma = (\Gamma_{103} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<b>7.9±1.2 OUR FIT</b>				
<b>7.3±1.2 OUR AVERAGE</b>				
7.4±0.8±1.1		<sup>1</sup> ARMS	05	CLE3 7.6 fb <sup>-1</sup> , E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
6.1±3.9±1.8		BARATE	98	ALEP 1991-1995 LEP runs
• • • We use the following data for averages but not for fits. • • •				
7.5±2.6±1.8		<sup>2</sup> RICHICHI	99	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<17	95	ABBIENDI	00D	OPAL 1990-1995 LEP runs

<sup>1</sup> Not independent of ARMS 05  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-\omega\nu_\tau)/\Gamma_{\text{total}}$  values.

<sup>2</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ ,  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ ,  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$  values.

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\eta))/\Gamma_{\text{total}} \quad \Gamma_{101}/\Gamma = (\Gamma_{103} + 0.892\Gamma_{177})/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID
<b>7.6±1.2 OUR FIT</b>	

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}} \quad \Gamma_{102}/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.7±0.5±0.8</b>	833	ARMS	05	CLE3 7.6 fb <sup>-1</sup> , E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega,\eta))/\Gamma_{\text{total}} \quad \Gamma_{103}/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID
<b>3.9±1.4 OUR FIT</b>	

$$\Gamma(K^-\pi^+K^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{104}/\Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.09</b>	95	BAUER	94	TPC E <sub>cm</sub> <sup>ee</sup> = 29 GeV

$$\Gamma(K^-\pi^+K^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{105}/\Gamma = (\Gamma_{106} + \Gamma_{107})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1496±0.0033 OUR FIT</b>				
<b>0.203 ±0.031 OUR AVERAGE</b>				
0.159 ±0.053 ±0.020		ABBIENDI	00D	OPAL 1990-1995 LEP runs
0.15 ±0.09 ±0.07 ±0.03	4	<sup>1</sup> BAUER	94	TPC E <sub>cm</sub> <sup>ee</sup> = 29 GeV

• • • We use the following data for averages but not for fits. • • •

0.238 ±0.042 <sup>2</sup> BARATE 98 ALEP 1991-1995 LEP runs

<sup>1</sup> We multiply 0.15% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$  values.

$$\Gamma(K^-\pi^+K^-\pi^-\nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{106}/\Gamma$$

VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.435±0.027 OUR FIT</b>				
<b>1.43 ±0.07 OUR AVERAGE</b>				

1.55 ±0.01 ±0.06 -0.05	108k	<sup>1</sup> LEE	10	BELL	666 fb <sup>-1</sup>	E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
1.346±0.010±0.036	18k	<sup>2</sup> AUBERT	08	BABR	342 fb <sup>-1</sup>	E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
1.55 ±0.06 ±0.09	932	<sup>3</sup> BRIERE	03	CLE3		E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
1.63 ±0.21 ±0.17		BARATE	98	ALEP		1991-1995 LEP runs
• • • We use the following data for averages but not for fits. • • •						
0.87 ±0.56 ±0.40		ABBIENDI	00D	OPAL		1990-1995 LEP runs
1.45 ±0.13 ±0.28	2.3k	<sup>4</sup> RICHICHI	99	CLEO		E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •						
2.2 <sup>+1.7</sup> <sub>-1.1</sub> ±0.5	9	<sup>5</sup> MILLS	85	DLCO		E <sub>cm</sub> <sup>ee</sup> = 29 GeV

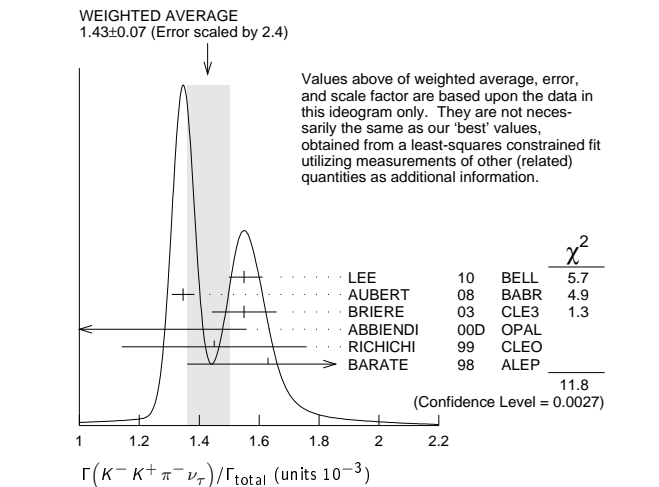
<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau)/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  value.

<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> 71% correlated with BRIERE 03  $\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  because of a common 5% normalization error.

<sup>4</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau)/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  and BALEST 95c  $\Gamma(\tau^- \rightarrow h^-h^+h^+\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$  values.

<sup>5</sup> Error correlated with MILLS 85 ( $K\pi\pi^0\nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.



$$\Gamma(K^-\pi^+K^-\pi^-\nu_\tau)/\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0)) \quad \Gamma_{106}/\Gamma_{68}$$

$$\Gamma_{106}/\Gamma_{68} = \Gamma_{106}/(\Gamma_{70} + 0.0153\Gamma_{176})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.592±0.030 OUR FIT</b>				
<b>1.83 ±0.05 OUR AVERAGE</b>				
1.60 ±0.15 ±0.30	2.3k	RICHICHI	99	CLEO E <sub>cm</sub> <sup>ee</sup> = 10.6 GeV

• • • We use the following data for averages but not for fits. • • •

1.84 ±0.01 ±0.05 108k <sup>1</sup> LEE 10 BELL 666 fb<sup>-1</sup> E<sub>cm</sub><sup>ee</sup> = 10.6 GeV

<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$  values.

$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{107}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.61 ± 0.18 OUR FIT</b>					
<b>0.60 ± 0.18 OUR AVERAGE</b>					
0.55 ± 0.14 ± 0.12	48		ARMS	05	CLE3 7.6 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
7.5 ± 2.9 ± 1.5			BARATE	98	ALEP 1991–1995 LEP runs
• • • We use the following data for averages but not for fits. • • •					
3.3 ± 1.8 ± 0.7	158		<sup>1</sup> RICHICHI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<27	95		ABBIENDI	00D	OPAL 1990–1995 LEP runs
<sup>1</sup> Not independent of RICHICHI 99					
$\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$ and BALEST 95c $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ values.					

$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau)/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau(\text{ex. } K^0))$					$\Gamma_{107}/\Gamma_{77}$
$\Gamma_{107}/\Gamma_{77} = \Gamma_{107}/(\Gamma_{78} + 0.892\Gamma_{176} + 0.0153\Gamma_{178})$					
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.14 ± 0.04 OUR FIT</b>					
<b>0.79 ± 0.44 ± 0.16</b>					
158		<sup>1</sup> RICHICHI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
<sup>1</sup> RICHICHI 99 also quote a 95%CL upper limit of 0.0157 for this measurement.					

$\Gamma(K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{108}/\Gamma = 0.489\Gamma_{168}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.2 ± 0.8 OUR FIT</b> Error includes scale factor of 5.4.					
<b>2.1 ± 0.8 OUR AVERAGE</b> Error includes scale factor of 5.4.					
3.29 ± 0.17 ± 0.19 − 0.20	3.2k		<sup>1</sup> LEE	10	BELL 666 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
1.58 ± 0.13 ± 0.12	275		<sup>2</sup> AUBERT	08	BABR 342 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 3.7	90		BRIERE	03	CLE3 $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
< 19	90		BARATE	98	ALEP 1991–1995 LEP runs
<sup>1</sup> See footnote to LEE 10 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements. Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$ value.					
<sup>2</sup> See footnote to AUBERT 08 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements.					

$\Gamma(K^- K^+ K^- \nu_\tau)/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$					$\Gamma_{108}/\Gamma_{68}$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.90 ± 0.02 ± 0.22 − 0.23	3.2k		<sup>1</sup> LEE	10	BELL 666 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
<sup>1</sup> Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ values.					

$\Gamma(K^- K^+ K^- \nu_\tau(\text{ex. } \phi))/\Gamma_{\text{total}}$					$\Gamma_{109}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 2.5 × 10<sup>-6</sup></b>					
90		AUBERT	08	BABR 342 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(K^- K^+ K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{110}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 4.8 × 10<sup>-6</sup></b>					
90		ARMS	05	CLE3 7.6 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{111}/\Gamma$
VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 0.25</b>					
95		BAUER	94	TPC $E_{\text{cm}}^{\text{ee}} = 29$ GeV	

$\Gamma(e^- e^- e^+ \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{112}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.8 ± 1.4 ± 0.4</b>					
5		ALAM	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{113}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 3.6</b>					
90		ALAM	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(3h^- 2h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^- \pi^+)(\text{"5-prong"}))/\Gamma_{\text{total}}$					$\Gamma_{114}/\Gamma$
$\Gamma_{114}/\Gamma = (\Gamma_{115} + \Gamma_{121})/\Gamma$					
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.099 ± 0.004 OUR FIT</b>					
<b>0.107 ± 0.007 OUR AVERAGE</b> Error includes scale factor of 1.1.					
0.170 ± 0.022 ± 0.026		<sup>1</sup> ACHARD	01D	L3 1992–1995 LEP runs	
0.097 ± 0.005 ± 0.011	419	GIBAUT	94B	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
0.102 ± 0.029	13	BYLSMA	87	HRS $E_{\text{cm}}^{\text{ee}} = 29$ GeV	
• • • We use the following data for averages but not for fits. • • •					
0.093 ± 0.009 ± 0.012		SCHAEEL	05c	ALEP 1991–1995 LEP runs	
0.115 ± 0.013 ± 0.006	112	<sup>2</sup> ABREU	01M	DLPH 1992–1995 LEP runs	
0.119 ± 0.013 ± 0.008	119	<sup>3</sup> ACKERSTAFF	99E	OPAL 1991–1995 LEP runs	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.26 ± 0.06 ± 0.05		ACTON	92H	OPAL $E_{\text{cm}}^{\text{ee}} = 88.2\text{--}94.2$ GeV	
0.10 ± 0.05 − 0.04		DECAMP	92c	ALEP 1989–1990 LEP runs	
0.16 ± 0.13 ± 0.04		BEHREND	89B	CELL $E_{\text{cm}}^{\text{ee}} = 14\text{--}47$ GeV	
0.3 ± 0.1 ± 0.2		BARTEL	85F	JADE $E_{\text{cm}}^{\text{ee}} = 34.6$ GeV	
0.13 ± 0.04	10	BELTRAMI	85	HRS Repl. by BYLSMA 87	
0.16 ± 0.08 ± 0.04	4	BURCHAT	85	MRK2 $E_{\text{cm}}^{\text{ee}} = 29$ GeV	
1.0 ± 0.4	10	BEHREND	82	CELL Repl. by BEHREND 89b	
<sup>1</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of $B(\tau^- \rightarrow \text{"1-prong"})$ and $B(\tau^- \rightarrow \text{"3-prong"})$ are $-0.082$ and $-0.19$ respectively.					
<sup>2</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of $B(\tau^- \rightarrow \text{1-prong})$ and $B(\tau^- \rightarrow \text{3-prong})$ are $-0.08$ and $-0.08$ respectively.					
<sup>3</sup> Not independent of ACKERSTAFF 99E $B(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau(\text{ex. } K^0))$ and $B(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau(\text{ex. } K^0))$ measurements.					

$\Gamma(3h^- 2h^+ \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$					$\Gamma_{115}/\Gamma = (\Gamma_{116} + \Gamma_{118} + 0.0153\Gamma_{183})/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>8.22 ± 0.32 OUR FIT</b>					
<b>8.32 ± 0.35 OUR AVERAGE</b>					
9.7 ± 1.5 ± 0.5	96	<sup>1</sup> ABDALLAH	06A	DLPH 1992–1995 LEP runs	
7.2 ± 0.9 ± 1.2	165	<sup>2</sup> SCHAEEL	05c	ALEP 1991–1995 LEP runs	
9.1 ± 1.4 ± 0.6	97	ACKERSTAFF	99E	OPAL 1991–1995 LEP runs	
7.7 ± 0.5 ± 0.9	295	GIBAUT	94B	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
6.4 ± 2.3 ± 1.0	12	ALBRECHT	88B	ARG $E_{\text{cm}}^{\text{ee}} = 10$ GeV	
5.1 ± 2.0	7	BYLSMA	87	HRS $E_{\text{cm}}^{\text{ee}} = 29$ GeV	
• • • We use the following data for averages but not for fits. • • •					
8.56 ± 0.05 ± 0.42	34k	AUBERT, B	05W	BABR 232 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
8.0 ± 1.1 ± 1.3	58	BUSKULIC	96	ALEP Repl. by SCHAEEL 05c	
6.7 ± 3.0	5	<sup>3</sup> BELTRAMI	85	HRS Repl. by BYLSMA 87	
<sup>1</sup> See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.					
<sup>2</sup> See footnote to SCHAEEL 05c $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.					
<sup>3</sup> The error quoted is statistical only.					

$\Gamma(3\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$					$\Gamma_{116}/\Gamma = (\Gamma_{117} + \Gamma_{171})/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b>8.21 ± 0.31 OUR FIT</b>					
• • • We use the following data for averages but not for fits. • • •					
<b>8.33 ± 0.04 ± 0.43</b>	<sup>1</sup> LEES	12X	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
<sup>1</sup> Not independent of LEES 12X $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau)/\Gamma$ and $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0, \omega, f_1(1285)))/\Gamma$ values.					

$\Gamma(3\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0, \omega, f_1(1285)))/\Gamma_{\text{total}}$					$\Gamma_{117}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>7.69 ± 0.30 OUR FIT</b>					
<b>7.68 ± 0.04 ± 0.40</b>	69k	LEES	12X	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(K^- 2\pi^- 2\pi^+ \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$					$\Gamma_{118}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b>0.6 ± 1.2 OUR FIT</b>					
<b>0.6 ± 0.5 ± 1.1</b>	<sup>1</sup> LEES	12X	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
<sup>1</sup> LEES 12X measurement corresponds to the lower limit of $< 2.4 \times 10^{-6}$ at 90% CL.					

$\Gamma(K^+ 3\pi^- \pi^+ \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{119}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 5.0 × 10<sup>-6</sup></b>					
90		LEES	12X	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(K^+ K^- 2\pi^- \pi^+ \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{120}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 4.5 × 10<sup>-7</sup></b>					
90		LEES	12X	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

## Lepton Particle Listings

 $\tau$  $\Gamma(3h^-2h^+\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$   $\Gamma_{121}/\Gamma = (\Gamma_{122}+\Gamma_{125})/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.64 \pm 0.11</math> OUR FIT</b>				
<b><math>1.74 \pm 0.27</math> OUR AVERAGE</b>				
$1.6 \pm 1.2 \pm 0.6$	13	<sup>1</sup> ABDALLAH 06A	06A	DLPH 1992-1995 LEP runs
$2.1 \pm 0.7 \pm 0.9$	95	<sup>2</sup> SCHAEEL 05C	05C	ALEP 1991-1995 LEP runs
$1.7 \pm 0.2 \pm 0.2$	231	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$2.7 \pm 1.8 \pm 0.9$	23	ACKERSTAFF 99E	OPAL	1991-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.8 \pm 0.7 \pm 1.2$	18	BUSKULIC 96	ALEP	Repl. by SCHAEEL 05C
$1.9 \pm 0.4 \pm 0.4$	31	GIBAUT 94B	CLEO	Repl. by ANASTASSOV 01
$5.1 \pm 2.2$	6	BYLSMA 87	HRS	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
$6.7 \pm 3.0$	5	<sup>3</sup> BELTRAMI 85	HRS	Repl. by BYLSMA 87

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> SCHAEEL 05C quote  $(1.4 \pm 0.7 \pm 0.9) \times 10^{-4}$ . We add  $0.7 \times 10^{-4}$  to remove their correction for  $\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  decays. See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> The error quoted is statistical only.

 $\Gamma(3\pi^-2\pi^+\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$   $\Gamma_{122}/\Gamma = (\Gamma_{124}+0.2292\Gamma_{158}+0.892\Gamma_{183})/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.62 \pm 0.11</math> OUR FIT</b>			
• • • We use the following data for averages but not for fits. • • •			
<b><math>1.65 \pm 0.05 \pm 0.09</math></b>	<sup>1</sup> LEES 12X	BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> Not independent of LEES 12X measurements of  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau(\text{ex.}K^0))/\Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.}K^0))/\Gamma$ , and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau(\text{ex.}K^0, \eta, \omega, f_1(1285)))/\Gamma$ .

 $\Gamma(3\pi^-2\pi^+\pi^0\nu_\tau(\text{ex.}K^0, \eta, f_1(1285)))/\Gamma_{\text{total}}$   $\Gamma_{123}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.11 \pm 0.04 \pm 0.09</math></b>	<sup>1</sup> LEES 12X	BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$1.11 \pm 0.04 \pm 0.09$	7.3k	LEES	12X BABR $468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau(\text{ex.}K^0))/\Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau(\text{ex.}K^0, \eta, \omega, f_1(1285)))/\Gamma$  values.

 $\Gamma(3\pi^-2\pi^+\pi^0\nu_\tau(\text{ex.}K^0, \eta, \omega, f_1(1285)))/\Gamma_{\text{total}}$   $\Gamma_{124}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.38 \pm 0.09</math> OUR FIT</b>				
<b><math>0.36 \pm 0.03 \pm 0.09</math></b>	7.3k	LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(K^-2\pi^-2\pi^+\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$   $\Gamma_{125}/\Gamma$ 

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 0.6</math> OUR FIT</b>			
<b><math>1.1 \pm 0.4 \pm 0.4</math></b>	<sup>1</sup> LEES 12X	BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$1.1 \pm 0.4 \pm 0.4$			

<sup>1</sup> LEES 12X measurement corresponds to the lower limit of  $< 1.9 \times 10^{-6}$  at 90% CL.

 $\Gamma(K^+3\pi^-\pi^+\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{126}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 8 \times 10^{-7}</math></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(3h^-2h^+2\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{127}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.4 \times 10^{-6}</math></b>	90	AUBERT,B	06 BABR	$232 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.1 \times 10^{-4}$	90	GIBAUT	94B CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma((5\pi^-)\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{128}/\Gamma$ 

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>0.78 \pm 0.05</math> OUR FIT</b>			
• • • We use the following data for averages but not for fits. • • •			
<b><math>0.61 \pm 0.06 \pm 0.08</math></b>	<sup>1</sup> GIBAUT 94B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

<sup>1</sup> Not independent of GIBAUT 94B  $B(3h^-2h^+\pi^0\nu_\tau)$ , PROCARIO 93  $B(h^-4\pi^0\nu_\tau)$ , and BORTOLETTO 93  $B(2h^-h^+2\pi^0\nu_\tau)/B(\text{"3prong"})$  measurements. Result is corrected for  $\eta$  contributions.

 $\Gamma(4h^-3h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{"7-prong"}))/\Gamma_{\text{total}}$   $\Gamma_{129}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.0 \times 10^{-7}</math></b>	90	AUBERT,B	05F BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.8 \times 10^{-5}$	95	ACKERSTAFF 97J	OPAL	1990-1995 LEP runs
$< 2.4 \times 10^{-6}$	90	EDWARDS 97B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$< 2.9 \times 10^{-4}$	90	BYLSMA 87	HRS	$E_{\text{cm}}^{\text{ee}} = 29$ GeV

 $\Gamma(4h^-3h^+\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{130}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.3 \times 10^{-7}</math></b>	90	AUBERT,B	05F BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(4h^-3h^+\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{131}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.5 \times 10^{-7}</math></b>	90	AUBERT,B	05F BABR	$232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(X^-(S=-1)\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{132}/\Gamma$ 

$$\Gamma_{132}/\Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{36} + \Gamma_{41} + \Gamma_{45} + \Gamma_{61} + \Gamma_{97} + \Gamma_{103} + \Gamma_{118} + \Gamma_{125} + \Gamma_{150} + \Gamma_{152} + \Gamma_{154} + 0.8312\Gamma_{168} + \Gamma_{177})/\Gamma$$

 $\Gamma(K^*(892)^- \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{133}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.92 \pm 0.04</math> OUR FIT</b>				
• • • We use the following data for averages but not for fits. • • •				
<b><math>2.87 \pm 0.12</math></b>	<sup>1</sup> BARATE 99R	ALEP	1991-1995 LEP runs	

<sup>1</sup> BARATE 99R perform a combined analysis of all ALEPH LEP 1 data on  $\tau$  branching fraction measurements for decay modes having total strangeness equal to  $-1$ .

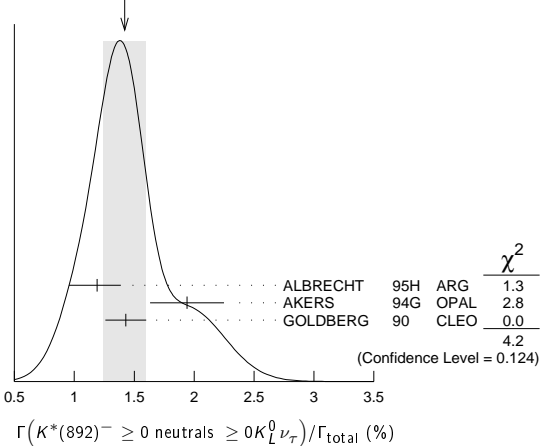
 $\Gamma(K^*(892)^- \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{133}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.42 \pm 0.18</math> OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
$1.19 \pm 0.15 \pm 0.13$	104	ALBRECHT 95H	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4$ -10.6 GeV
$1.94 \pm 0.27 \pm 0.15$	74	<sup>1</sup> AKERS 94G	OPAL	$E_{\text{cm}}^{\text{ee}} = 88$ -94 GeV
$1.43 \pm 0.11 \pm 0.13$	475	<sup>2</sup> GOLDBERG 90	CLEO	$E_{\text{cm}}^{\text{ee}} = 9.4$ -10.9 GeV

<sup>1</sup> AKERS 94G reject events in which a  $K_S^0$  accompanies the  $K^*(892)^-$ . We do not correct for them.

<sup>2</sup> GOLDBERG 90 estimates that 10% of observed  $K^*(892)$  are accompanied by a  $\pi^0$ .

WEIGHTED AVERAGE  
 $1.42 \pm 0.18$  (Error scaled by 1.4)

 $\Gamma(K^*(892)^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{134}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.20 \pm 0.07</math> OUR AVERAGE</b>				Error includes scale factor of 1.8. See the ideogram below.
$1.131 \pm 0.006 \pm 0.051$	49k	<sup>1</sup> EPIFANOV 07	BELL	$351 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$1.326 \pm 0.063$		BARATE 99R	ALEP	1991-1995 LEP runs
$1.11 \pm 0.12$		<sup>2</sup> COAN 96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV
$1.42 \pm 0.22 \pm 0.09$		<sup>3</sup> ACCIARRI 95F	L3	1991-1993 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.39 \pm 0.09 \pm 0.10$		<sup>4</sup> BUSKULIC 96	ALEP	Repl. by BARATE 99R
$1.45 \pm 0.13 \pm 0.11$	273	<sup>5</sup> BUSKULIC 94F	ALEP	Repl. by BUSKULIC 96
$1.23 \pm 0.21 \pm 0.11$	54	<sup>6</sup> ALBRECHT 88L	ARG	$E_{\text{cm}}^{\text{ee}} = 10$ GeV
$1.9 \pm 0.3 \pm 0.4$	44	<sup>7</sup> TSCHIRHART 88	HRS	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
$1.5 \pm 0.4 \pm 0.4$	15	<sup>8</sup> AIHARA 87C	TPC	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
$1.3 \pm 0.3 \pm 0.3$	31	YELTON 86	MRK2	$E_{\text{cm}}^{\text{ee}} = 29$ GeV
$1.7 \pm 0.7$	11	DORFAN 81	MRK2	$E_{\text{cm}}^{\text{ee}} = 4.2$ -6.7 GeV

<sup>1</sup> EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) = (3.77 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \pm 0.12(\text{mod})) \times 10^{-3}$ . We add the systematic and model uncertainties in quadrature and divide by  $B(K^*(892)^- \rightarrow K_S^0 \pi^-) = 0.3333$ .

<sup>2</sup> Not independent of COAN 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and BATTLE 94  $B(K^- \pi^0 \nu_\tau)$  measurements.  $K$  final states are consistent with and assumed to originate from  $K^*(892)^-$  production.

<sup>3</sup> This result is obtained from their  $B(\pi^- \bar{K}^0 \nu_\tau)$  assuming all those decays originate in  $K^*(892)^-$  decays.

<sup>4</sup> Not independent of BUSKULIC 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and  $B(K^- \pi^0 \nu_\tau)$  measurements.

<sup>5</sup> BUSKULIC 94F obtain this result from BUSKULIC 94F  $B(\bar{K}^0 \pi^- \nu_\tau)$  and BUSKULIC 94E  $B(K^- \pi^0 \nu_\tau)$  assuming all of those decays originate in  $K^*(892)^-$  decays.

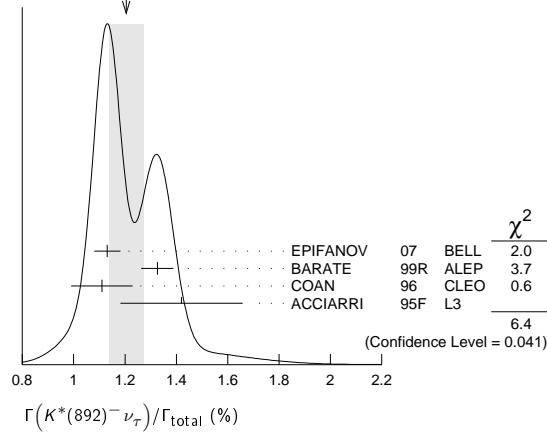
<sup>6</sup> The authors divide by  $\Gamma_2/\Gamma = 0.865$  to obtain this result.

<sup>7</sup> Not independent of TSCHIRHART 88  $\Gamma(\tau^- \rightarrow h^- \bar{K}^0 \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau) / \Gamma$ .

<sup>8</sup> Decay  $\pi^-$  identified in this experiment, is assumed in the others.



WEIGHTED AVERAGE  
1.20±0.07 (Error scaled by 1.8)



### $\Gamma(K^*(892)^-\nu_\tau)/\Gamma(\pi^-\pi^0\nu_\tau)$ $\Gamma_{134}/\Gamma_{14}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.075 ± 0.027</b>	<sup>1</sup> ABREU 94K	DLPH	LEP 1992 Z data

<sup>1</sup> ABREU 94K quote  $B(\tau^- \rightarrow K^*(892)^-\nu_\tau)B(K^*(892)^- \rightarrow K^-\pi^0)/B(\tau^- \rightarrow \rho^-\nu_\tau) = 0.025 \pm 0.009$ . We divide by  $B(K^*(892)^- \rightarrow K^-\pi^0) = 0.333$  to obtain this result.

### $\Gamma(K^*(892)^-\nu_\tau \rightarrow \pi^-\bar{K}^0\nu_\tau)/\Gamma(\pi^-\bar{K}^0\nu_\tau)$ $\Gamma_{135}/\Gamma_{36}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.933 ± 0.027</b>	49k	EPIFANOV 07	BELL	351 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV

### $\Gamma(K^*(892)^0 K^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{total}$ $\Gamma_{136}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.32 ± 0.08 ± 0.12</b>	119	GOLDBERG 90	CLEO	$E_{cm}^{ee} = 9.4\text{--}10.9$ GeV

### $\Gamma(K^*(892)^0 K^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{137}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21 ± 0.04 OUR AVERAGE</b>				
0.213 ± 0.048		<sup>1</sup> BARATE 98	ALEP	1991–1995 LEP runs
0.20 ± 0.05 ± 0.04	47	ALBRECHT 95H	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup> BARATE 98 measure the  $K^-(\rho^0 \rightarrow \pi^+\pi^-)$  fraction in  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  decays to be  $(35 \pm 11)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma_{total}$  assuming the intermediate states are all  $K^-\rho$  and  $K^-\bar{K}^*(892)^0$ .

### $\Gamma(\bar{K}^*(892)^0 \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{total}$ $\Gamma_{138}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.38 ± 0.11 ± 0.13</b>	105	GOLDBERG 90	CLEO	$E_{cm}^{ee} = 9.4\text{--}10.9$ GeV

### $\Gamma(\bar{K}^*(892)^0 \pi^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{139}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22 ± 0.05 OUR AVERAGE</b>				
0.209 ± 0.058		<sup>1</sup> BARATE 98	ALEP	1991–1995 LEP runs
0.25 ± 0.10 ± 0.05	27	ALBRECHT 95H	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup> BARATE 98 measure the  $K^-\bar{K}^*(892)^0$  fraction in  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  decays to be  $(87 \pm 13)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma_{total}$ .

### $\Gamma((\bar{K}^*(892)^0 \pi^-) \nu_\tau \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau)/\Gamma_{total}$ $\Gamma_{140}/\Gamma$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.10 ± 0.04 OUR AVERAGE</b>			
0.097 ± 0.044 ± 0.036	<sup>1</sup> BARATE 99K	ALEP	1991–1995 LEP runs
0.106 ± 0.037 ± 0.032	<sup>2</sup> BARATE 98E	ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_S^0$ 's in their hadron calorimeter. They determine the  $\bar{K}^0\rho^-$  fraction in  $\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$  decays to be  $(0.72 \pm 0.12 \pm 0.10)$  and multiply their  $B(\pi^-\bar{K}^0\pi^0\nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

<sup>2</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+\pi^-$  decays. They determine the  $\bar{K}^0\rho^-$  fraction in  $\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$  decays to be  $(0.64 \pm 0.09 \pm 0.10)$  and multiply their  $B(\pi^-\bar{K}^0\pi^0\nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

### $\Gamma(K_1(1270)^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{141}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.47 ± 0.11 OUR AVERAGE</b>				
0.48 ± 0.11		BARATE 99R	ALEP	1991–1995 LEP runs
0.41 ± 0.41 ± 0.10	5	<sup>1</sup> BAUER 94	TPC	$E_{cm}^{ee} = 29$ GeV

<sup>1</sup> We multiply 0.41% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

### $\Gamma(K_1(1400)^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{142}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.17 ± 0.26 OUR AVERAGE</b>				Error includes scale factor of 1.7.
0.05 ± 0.17		BARATE 99R	ALEP	1991–1995 LEP runs
0.76 ± 0.40 ± 0.20	11	<sup>1</sup> BAUER 94	TPC	$E_{cm}^{ee} = 29$ GeV

<sup>1</sup> We multiply 0.76% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

### $[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]/\Gamma_{total}$ $(\Gamma_{141} + \Gamma_{142})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.17 ± 0.41 ± 0.37 ± 0.29</b>	16	<sup>1</sup> BAUER 94	TPC	$E_{cm}^{ee} = 29$ GeV

<sup>1</sup> We multiply 1.17% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error. Not independent of BAUER 94  $B(K_1(1270)^-\nu_\tau)$  and BAUER 94  $B(K_1(1400)^-\nu_\tau)$  measurements.

### $\Gamma(K_1(1270)^-\nu_\tau)/[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]$ $\Gamma_{141}/(\Gamma_{141} + \Gamma_{142})$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.69 ± 0.15 OUR AVERAGE</b>			
0.71 ± 0.16 ± 0.11	<sup>1</sup> ABBIENDI 00D	OPAL	1990–1995 LEP runs
0.66 ± 0.19 ± 0.13	<sup>2</sup> ASNER 00B	CLEO	$E_{cm}^{ee} = 10.6$ GeV

<sup>1</sup> ABBIENDI 00D assume the resonance structure of  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  decays is dominated by the  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

<sup>2</sup> ASNER 00B assume the resonance structure of  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  (ex.  $K^0$ ) decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

### $\Gamma(K^*(1410)^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{143}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>1.5 ± 1.4</b>	BARATE 99R	ALEP	1991–1995 LEP runs

### $\Gamma(K_0^*(1430)^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{144}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.5</b>	95	BARATE 99R	ALEP	1991–1995 LEP runs

### $\Gamma(K_2^*(1430)^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{145}/\Gamma$

VALUE (%)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.3</b>	95		TSCHIRHART 88	HRS	$E_{cm}^{ee} = 29$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.33	95		<sup>1</sup> ACCIARRI 95F	L3	1991–1993 LEP runs
< 0.9	95	0	DORFAN 81	MRK2	$E_{cm}^{ee} = 4.2\text{--}6.7$ GeV

<sup>1</sup> ACCIARRI 95F quote  $B(\tau^- \rightarrow K^*(1430)^- \rightarrow \pi^-\bar{K}^0\nu_\tau) < 0.11\%$ . We divide by  $B(K^*(1430)^- \rightarrow \pi^-\bar{K}^0) = 0.33$  to obtain the limit shown.

### $\Gamma(a_0(980)^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{total} \times B(a_0(980)^- \rightarrow K^0 K^-)$ $\Gamma_{146}/\Gamma \times B$

VALUE (units 10 <sup>-4</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.8</b>	90	GOLDBERG 90	CLEO	$E_{cm}^{ee} = 9.4\text{--}10.9$ GeV

### $\Gamma(\eta\pi^-\nu_\tau)/\Gamma_{total}$ $\Gamma_{147}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.99</b>	95		<sup>1</sup> DEL-AMO-SA...11E	BABR	470 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 6.2	95		BUSKULIC	97c	ALEP	1991–1994 LEP runs
< 1.4	95	0	BARTELT	96	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
< 3.4	95		ARTUSO	92	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
< 90	95		ALBRECHT	88M	ARG	$E_{\text{cm}}^{ee} \approx 10$ GeV
<140	90		BEHREND	88	CELL	$E_{\text{cm}}^{ee} = 14\text{--}46.8$ GeV
<180	95		BARINGER	87	CLEO	$E_{\text{cm}}^{ee} = 10.5$ GeV
<250	90	0	COFFMAN	87	MRK3	$E_{\text{cm}}^{ee} = 3.77$ GeV
510 ±100±120		65	DERRICK	87	HRS	$E_{\text{cm}}^{ee} = 29$ GeV
<100	95		GAN	87B	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>1</sup> DEL-AMO-SANCHEZ 11E also quote  $B(\tau^- \rightarrow \eta\pi^-\nu_\tau) = (3.4 \pm 3.4 \pm 2.1) \times 10^{-5}$ .

### $\Gamma(\eta\pi^-\pi^0\nu_\tau)/\Gamma_{total}$ $\Gamma_{148}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.39 ± 0.07 OUR FIT</b>					
<b>1.38 ± 0.09 OUR AVERAGE</b>					Error includes scale factor of 1.2.
1.35 ± 0.03 ± 0.07	6.0k		INAMI 09	BELL	490 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV

1.8 ± 0.4 ± 0.2

1.7 ± 0.2 ± 0.2

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BUSKULIC 97c ALEP 1991–1994 LEP runs

ARTUSO 92 CLEO  $E_{cm}^{ee} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 11.0	95	ALBRECHT 88M	ARG	$E_{cm}^{ee} \approx 10$ GeV
< 21.0	95	BARINGER 87	CLEO	$E_{cm}^{ee} = 10.5$ GeV
42.0 ± 7.0 ± 12.0 ± 16.0		<sup>1</sup> GAN 87	MRK2	$E_{cm}^{ee} = 29$ GeV

<sup>1</sup> Highly correlated with GAN 87  $\Gamma(\pi^- 3\pi^0\nu_\tau)/\Gamma_{total}$  value.

# Lepton Particle Listings

$\tau$

$\Gamma(\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{149}/\Gamma$
VALUE (units $10^{-4}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.9 ± 0.4 OUR FIT</b>					
<b>1.81 ± 0.31 OUR AVERAGE</b>					
2.01 ± 0.34 ± 0.22	381	LEES	12x BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

- • • We use the following data for averages but not for fits. • • •
- 1.5 ± 0.5 30 <sup>1</sup>ANASTASSOV 01 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV
- • • We do not use the following data for averages, fits, limits, etc. • • •
- 1.4 ± 0.6 ± 0.3 15 <sup>2</sup>BERGFELD 97 CLEO Repl. by ANAS-TASSOV 01
- < 4.3 95 ARTUSO 92 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV
- < 120 95 ALBRECHT 88M ARG  $E_{\text{cm}}^{\text{ee}} \approx 10$  GeV

<sup>1</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 value of  $(1.5 \pm 0.6 \pm 0.3) \times 10^{-4}$  obtained using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays.  
<sup>2</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  decays.

$\Gamma(\eta K^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{150}/\Gamma$
VALUE (units $10^{-4}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.55 ± 0.08 OUR FIT</b>					
<b>1.54 ± 0.08 OUR AVERAGE</b>					
1.42 ± 0.11 ± 0.07	690	DEL-AMO-SA..11E	BABR	470 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
1.58 ± 0.05 ± 0.09	1.6k	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
2.9 <sup>+1.3</sup> <sub>-1.2</sub> ± 0.7		BUSKULIC 97c	ALEP	1991-1994 LEP runs	
2.6 ± 0.5 ± 0.5	85	BARTELT 96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV	
< 4.7	95	ARTUSO 92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV	

$\Gamma(\eta K^*(892)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{151}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.38 ± 0.15 OUR AVERAGE</b>					
1.34 ± 0.12 ± 0.09	245	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
2.90 ± 0.80 ± 0.42	25	BISHAI 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

<sup>1</sup> Not independent of INAMI 09  $B(\tau^- \rightarrow \eta K^-\pi^0\nu_\tau)$  and  $B(\tau^- \rightarrow \eta \bar{K}^0\pi^-\nu_\tau)$  values.

$\Gamma(\eta K^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{152}/\Gamma$
VALUE (units $10^{-4}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.48 ± 0.12 OUR FIT</b>					
<b>0.48 ± 0.12 OUR AVERAGE</b>					
0.46 ± 0.11 ± 0.04	270	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
1.77 ± 0.56 ± 0.71	36	BISHAI 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\eta K^-\pi^0(\text{non-}K^*(892))\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{153}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 3.5 × 10<sup>-5</sup></b>	90	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\eta \bar{K}^0\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{154}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.94 ± 0.15 OUR FIT</b>					
<b>0.93 ± 0.15 OUR AVERAGE</b>					
0.88 ± 0.14 ± 0.06	161	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
2.20 ± 0.70 ± 0.22	15	<sup>2</sup> BISHAI 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0\pi^-\nu_\tau) = (0.44 \pm 0.07 \pm 0.03) \times 10^{-4}$  by 2 to obtain the listed value.  
<sup>2</sup> We multiply the BISHAI 99 measurement  $B(\tau^- \rightarrow \eta K_S^0\pi^-\nu_\tau) = (1.10 \pm 0.35 \pm 0.11) \times 10^{-4}$  by 2 to obtain the listed value.

$\Gamma(\eta \bar{K}^0\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{155}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 5.0 × 10<sup>-5</sup></b>	90	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0\pi^-\pi^0\nu_\tau) < 2.5 \times 10^{-5}$  by 2 to obtain the listed value.

$\Gamma(\eta K^-\bar{K}^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{156}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 9.0 × 10<sup>-6</sup></b>	90	<sup>1</sup> INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K^-\bar{K}_S^0\nu_\tau) < 4.5 \times 10^{-6}$  by 2 to obtain the listed value.

$\Gamma(\eta\pi^+\pi^-\pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{157}/\Gamma$
VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 0.3</b>	90	ABACHI 87B	HRS	$E_{\text{cm}}^{\text{ee}} = 29$ GeV	

$\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$					$\Gamma_{158}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.19 ± 0.13 OUR FIT</b>					
<b>2.23 ± 0.12 OUR AVERAGE</b>					
2.10 ± 0.09 ± 0.13	2.9k	<sup>1</sup> LEES	12x BABR	$\eta \rightarrow \gamma\gamma$	
2.37 ± 0.12 ± 0.18	1.4k	<sup>1</sup> LEES	12x BABR	$\eta \rightarrow \pi^+\pi^-\pi^0$	
2.54 ± 0.27 ± 0.25	315	<sup>1</sup> LEES	12x BABR	$\eta \rightarrow 3\pi^0$	
2.3 ± 0.5	170	<sup>2</sup> ANASTASSOV 01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
1.60 ± 0.05 ± 0.11	1.8 k	AUBERT	08AE BABR	Repl. by LEES 12x	
3.4 <sup>+0.6</sup> <sub>-0.5</sub> ± 0.6	89	<sup>3</sup> BERGFELD 97	CLEO	Repl. by ANASTASSOV 01	

<sup>1</sup> LEES 12x uses 468 fb<sup>-1</sup> of data taken at  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV. It gives the average of the three measurements listed here as  $(2.25 \pm 0.07 \pm 0.12) \times 10^{-4}$ .  
<sup>2</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 measurements using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+\pi^-\pi^0$  and  $\eta \rightarrow 3\pi^0$  decays.  
<sup>3</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow 3\pi^0$  decays.

$\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0,f_1(1285)))/\Gamma_{\text{total}}$					$\Gamma_{159}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.99 ± 0.09 ± 0.13</b>					
		<sup>1</sup> LEES	12x BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

<sup>1</sup> LEES 12x obtain this result by subtracting their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement from their  $B(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  measurement.

$\Gamma(\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{160}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 3.9 × 10<sup>-4</sup></b>	90	BERGFELD 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\eta\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{161}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 7.4 × 10<sup>-6</sup></b>	90	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
< 1.1 × 10 <sup>-4</sup>	95	ARTUSO 92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV	
< 8.3 × 10 <sup>-3</sup>	95	ALBRECHT 88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10$ GeV	

$\Gamma(\eta\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{162}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 2.0</b>	95	ARTUSO 92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV	
< 90	95	ALBRECHT 88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10$ GeV	

$\Gamma(\eta\eta K^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{163}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 3.0 × 10<sup>-6</sup></b>	90	INAMI 09	BELL	490 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\eta'(958)\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{164}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 4.0 × 10<sup>-6</sup></b>	90	LEES	12x BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
< 7.2 × 10 <sup>-6</sup>	90	AUBERT	08AE BABR	384 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
< 7.4 × 10 <sup>-5</sup>	90	BERGFELD 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\eta'(958)\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{165}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 1.2 × 10<sup>-5</sup></b>	90	LEES	12x BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
< 8.0 × 10 <sup>-5</sup>	90	BERGFELD 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\eta'(958)K^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{166}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 2.4 × 10<sup>-6</sup></b>	90	LEES	12x BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	

$\Gamma(\phi\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{167}/\Gamma$
VALUE (units $10^{-5}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3.42 ± 0.55 ± 0.25</b>	344	AUBERT	08 BABR	342 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
< 20	90	<sup>1</sup> AVERY 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV	
< 35	90	ALBRECHT 95H	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV	

<sup>1</sup> AVERY 97 limit varies from  $(1.2\text{--}2.0) \times 10^{-4}$  depending on decay model assumptions.

$\Gamma(\phi K^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{168}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**4.4 ±1.6 OUR FIT****3.70±0.33 OUR AVERAGE** Error includes scale factor of 1.3.

• • • We use the following data for averages but not for fits. • • •

3.39±0.20±0.28	274	AUBERT	08	BABR	342 fb <sup>-1</sup> $E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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4.05±0.25±0.26	551	INAMI	06	BELL	401 fb <sup>-1</sup> $E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<6.7	90	<sup>1</sup> AVERY	97	CLEO	$E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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<sup>1</sup> AVERY 97 limit varies from  $(5.4-6.7) \times 10^{-5}$  depending on decay model assumptions. $\Gamma(\hat{f}_1(1285)\pi^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{169}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**3.9 ±0.5 OUR AVERAGE** Error includes scale factor of 1.9.

4.73±0.28±0.45	3.7k	<sup>1</sup> LEES	12x	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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3.60±0.18±0.23	2.5k	<sup>2</sup> LEES	12x	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

3.19±0.18±1.00	1.3 k	<sup>3</sup> AUBERT	08AE	BABR	Repl. by LEES 12x
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3.9 ±0.7 ±0.5	1.4 k	<sup>4</sup> AUBERT,B	05W	BABR	Repl. by LEES 12x
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5.8 <sup>+1.4</sup> <sub>-1.3</sub> ±1.8	54	<sup>5</sup> BERGFELD	97	CLEO	$E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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<sup>1</sup> LEES 12x obtain this value by dividing their  $B(\tau^- \rightarrow \hat{f}_1(1285)\pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau)$  measurement by the PDG 12 value of  $B(\hat{f}_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.111 \pm_{-0.006}^{+0.007}$ .<sup>2</sup> LEES 12x obtain this value by dividing their  $B(\tau^- \rightarrow \hat{f}_1(1285)\pi^- \nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)$  measurement by 2/3 of the PDG 12 value of  $B(\hat{f}_1(1285) \rightarrow \eta\pi\pi) = 0.524 \pm_{-0.021}^{+0.019}$ .<sup>3</sup> AUBERT 08AE obtain this value by dividing their  $B(\tau^- \rightarrow \hat{f}_1(1285)\pi^- \nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)$  measurement by the PDG 06 value of  $B(\hat{f}_1(1285) \rightarrow \eta\pi^- \pi^+) = 0.35 \pm 0.11$ . The quote  $(3.19 \pm 0.18 \pm 0.16 \pm 0.99) \times 10^{-4}$  where the final error is due to the uncertainty on  $B(\hat{f}_1(1285) \rightarrow \eta\pi^- \pi^+)$ . We combine the two systematic errors in quadrature.<sup>4</sup> AUBERT,B 05W use the  $\hat{f}_1(1285) \rightarrow 2\pi^+ 2\pi^-$  decay mode and the PDG 04 value of  $B(\hat{f}_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.110 \pm_{-0.006}^{+0.007}$ .<sup>5</sup> BERGFELD 97 use the  $\hat{f}_1(1285) \rightarrow \eta\pi^+ \pi^-$  decay mode. $\Gamma(\hat{f}_1(1285)\pi^- \nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{170}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.18±0.07 OUR AVERAGE** Error includes scale factor of 1.3.

1.26±0.06±0.06	2.5k	LEES	12x	BABR	468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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1.11±0.06±0.05	1.3 k	AUBERT	08AE	BABR	384 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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 $\Gamma(\hat{f}_1(1285)\pi^- \nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)/\Gamma(\eta\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$   $\Gamma_{170}/\Gamma_{158}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.69±0.01±0.05** <sup>1</sup> AUBERT 08AE BABR 384 fb<sup>-1</sup>,  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.55±0.14	BERGFELD	97	CLEO	$E_{\text{cm}}^{\text{e}} = 10.6$ GeV
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<sup>1</sup> Not independent of AUBERT 08AE  $B(\tau^- \rightarrow \hat{f}_1(1285)\pi^- \nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)$  and  $B(\tau^- \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$  values. $\Gamma(\hat{f}_1(1285)\pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{171}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.52 ±0.04 OUR FIT****0.520±0.031±0.037** 3.7k LEES 12x BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV $\Gamma(\pi(1300)^- \nu_\tau \rightarrow (\rho\pi)^- \nu_\tau \rightarrow (3\pi)^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{172}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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<1.0 × 10<sup>-4</sup> 90 ASNER 00 CLEO  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV $\Gamma(\pi(1300)^- \nu_\tau \rightarrow ((\pi\pi)_{\text{S-wave}} \pi)^- \nu_\tau \rightarrow (3\pi)^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{173}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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<1.9 × 10<sup>-4</sup> 90 ASNER 00 CLEO  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV $\Gamma(h^- \omega \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{174}/\Gamma$  $\Gamma_{174}/\Gamma = (\Gamma_{176} + \Gamma_{177} + \Gamma_{178})/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**2.40±0.08 OUR FIT**

• • • We use the following data for averages but not for fits. • • •

1.65±0.3 ±0.2	1513	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{e}} \approx 10$ GeV
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 $\Gamma(h^- \omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{175}/\Gamma = (\Gamma_{176} + \Gamma_{177})/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.99±0.06 OUR FIT****1.92±0.07 OUR AVERAGE**

1.91±0.07±0.06	5803	BUSKULIC	97c	ALEP	1991–1994 LEP runs
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1.60±0.27±0.41	139	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{e}} = 10.5$ GeV
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• • • We use the following data for averages but not for fits. • • •

1.95±0.07±0.11	2223	<sup>1</sup> BALEST	95c	CLEO	$E_{\text{cm}}^{\text{e}} \approx 10.6$ GeV
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<sup>1</sup> Not independent of BALEST 95c  $B(\tau^- \rightarrow h^- \omega \nu_\tau)/B(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau)$  value. $[\Gamma(\pi^- \omega \nu_\tau) + \Gamma(K^- \omega \nu_\tau)]/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau(\text{ex. } K^0))$   $(\Gamma_{176} + \Gamma_{177})/\Gamma_{74}$  $(\Gamma_{176} + \Gamma_{177})/\Gamma_{74} = (\Gamma_{176} + \Gamma_{177})/(\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{176} + 0.892\Gamma_{177} + 0.0153\Gamma_{178})$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**43.5±1.4 OUR FIT****45.3±1.9 OUR AVERAGE**

43.1±3.3	2350	<sup>1</sup> BUSKULIC	96	ALEP	LEP 1991–1993 data
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46.4±1.6±1.7	2223	<sup>2</sup> BALEST	95c	CLEO	$E_{\text{cm}}^{\text{e}} \approx 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

37 ±5 ±2	458	<sup>3</sup> ALBRECHT	91D	ARG	$E_{\text{cm}}^{\text{e}} = 9.4-10.6$ GeV
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<sup>1</sup> BUSKULIC 96 quote the fraction of  $\tau \rightarrow h^- h^- h^+ \pi^0 \nu_\tau(\text{ex. } K^0)$  decays which originate in a  $h^- \omega$  final state =  $0.383 \pm 0.029$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).<sup>2</sup> BALEST 95c quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau(\text{ex. } K^0)$  decays which originate in a  $h^- \omega$  final state equals  $0.412 \pm 0.014 \pm 0.015$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).<sup>3</sup> ALBRECHT 91D quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  decays which originate in a  $\pi^- \omega$  final state equals  $0.33 \pm 0.04 \pm 0.02$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888). $\Gamma(\pi^- \omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{176}/\Gamma$ 

VALUE (%)	DOCUMENT ID
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**1.95±0.06 OUR FIT** $\Gamma(K^- \omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{177}/\Gamma$ 

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VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**4.1 ±0.9 OUR FIT****4.1±0.6±0.7** 500 ARMS 05 CLE3 7.6 fb<sup>-1</sup>,  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV $\Gamma(h^- \omega \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{178}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.41±0.04 OUR FIT****0.43±0.06±0.05** 7283 BUSKULIC 97c ALEP 1991–1994 LEP runs $\Gamma(h^- \omega \pi^0 \nu_\tau)/\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)$   $\Gamma_{178}/\Gamma_{62}$  $\Gamma_{178}/\Gamma_{62} = \Gamma_{178}/(0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2292\Gamma_{149} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.3759\Gamma_{154} + 0.3268\Gamma_{158} + 0.7259\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178} + 0.892\Gamma_{180})$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**(2.69 ±0.28) × 10<sup>-2</sup> OUR FIT**

• • • We use the following data for averages but not for fits. • • •

**0.028±0.003±0.003** 430 <sup>1</sup> BORTOLETTO 93 CLEO  $E_{\text{cm}}^{\text{e}} \approx 10.6$  GeV<sup>1</sup> Not independent of BORTOLETTO 93  $\Gamma(\tau^- \rightarrow h^- \omega \pi^0 \nu_\tau)/\Gamma(\tau^- \rightarrow h^- h^- h^+ 2\pi^0 \nu_\tau(\text{ex. } K^0))$  value. $\Gamma(h^- \omega \pi^0 \nu_\tau)/\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau(\text{ex. } K^0))$   $\Gamma_{178}/\Gamma_{84}$  $\Gamma_{178}/\Gamma_{84} = \Gamma_{178}/(\Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
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**82±8 OUR FIT****81±6±6** BORTOLETTO 93 CLEO  $E_{\text{cm}}^{\text{e}} \approx 10.6$  GeV $\Gamma(h^- \omega 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{179}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.4 ±0.4 ±0.3** 53 ANASTASSOV 01 CLEO  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.89 <sup>+0.74</sup> <sub>-0.67</sub> ±0.40	19	ANDERSON	97	CLEO	Repl. by ANASTASSOV 01
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 $\Gamma(\pi^- \omega 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{180}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.71±0.16 OUR FIT****0.73±0.12±0.12** 1.1k LEES 12x BABR 468 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV $\Gamma(h^- 2\omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{181}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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<5.4 × 10<sup>-7</sup> 90 AUBERT,B 06 BABR 232 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV $\Gamma(2h^- h^+ \omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{182}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.2±0.2±0.1** 110 ANASTASSOV 01 CLEO  $E_{\text{cm}}^{\text{e}} = 10.6$  GeV

# Lepton Particle Listings

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$\Gamma(2\pi^-\pi^+\omega\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$					$\Gamma_{183}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b><math>0.84\pm0.06</math> OUR FIT</b>					
<b><math>0.84\pm0.04\pm0.06</math></b>	2.4k	LEES	12x	BABR 468 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	

$\Gamma(e^-\gamma)/\Gamma_{\text{total}}$					$\Gamma_{184}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;3.3\times10^{-8}</math></b>	90	AUBERT	10B	BABR 516 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.2\times10^{-7}$	90	HAYASAKA	08	BELL 535 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.1\times10^{-7}$	90	AUBERT	06c	BABR 232 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.9\times10^{-7}$	90	HAYASAKA	05	BELL 86.7 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<2.7\times10^{-6}$	90	EDWARDS	97	CLEO	
$<1.1\times10^{-4}$	90	ABREU	95U	DLPH 1990–1993 LEP runs	
$<1.2\times10^{-4}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<2.0\times10^{-4}$	90	KEH	88	CBAL $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<6.4\times10^{-4}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

$\Gamma(\mu^-\gamma)/\Gamma_{\text{total}}$					$\Gamma_{185}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;4.4\times10^{-8}</math></b>	90	AUBERT	10B	BABR 516 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<4.5\times10^{-8}$	90	HAYASAKA	08	BELL 535 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.8\times10^{-8}$	90	AUBERT,B	05A	BABR 232 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.1\times10^{-7}$	90	ABE	04B	BELL 86.3 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.1\times10^{-6}$	90	AHMED	00	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.0\times10^{-6}$	90	EDWARDS	97	CLEO	
$<6.2\times10^{-5}$	90	ABREU	95U	DLPH 1990–1993 LEP runs	
$<0.42\times10^{-5}$	90	BEAN	93	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.4\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<55\times10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

$\Gamma(e^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{186}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;8.0\times10^{-8}</math></b>	90	MIYAZAKI	07	BELL 401 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.3\times10^{-7}$	90	AUBERT	07I	BABR 339 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.9\times10^{-7}$	90	ENARI	05	BELL 154 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.7\times10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<17\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<14\times10^{-5}$	90	KEH	88	CBAL $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<210\times10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

$\Gamma(\mu^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{187}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.1\times10^{-7}</math></b>	90	AUBERT	07I	BABR 339 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.2\times10^{-7}$	90	MIYAZAKI	07	BELL 401 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<4.1\times10^{-7}$	90	ENARI	05	BELL 154 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<4.0\times10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<4.4\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<82\times10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

$\Gamma(e^-K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{188}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2.6\times10^{-8}</math></b>	90	MIYAZAKI	10A	BELL 671 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<3.3\times10^{-8}$	90	AUBERT	09D	BABR 469 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<5.6\times10^{-8}$	90	MIYAZAKI	06A	BELL 281 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<9.1\times10^{-7}$	90	CHEN	02c	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.3\times10^{-3}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

$\Gamma(\mu^-K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{189}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2.3\times10^{-8}</math></b>	90	MIYAZAKI	10A	BELL 671 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<4.0\times10^{-8}$	90	AUBERT	09D	BABR 469 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<4.9\times10^{-8}$	90	MIYAZAKI	06A	BELL 281 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<9.5\times10^{-7}$	90	CHEN	02c	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.0\times10^{-3}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

$\Gamma(e^-\eta)/\Gamma_{\text{total}}$					$\Gamma_{190}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;9.2\times10^{-8}</math></b>	90	MIYAZAKI	07	BELL 401 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.6\times10^{-7}$	90	AUBERT	07I	BABR 339 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<2.4\times10^{-7}$	90	ENARI	05	BELL 154 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<8.2\times10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.3\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<24\times10^{-5}$	90	KEH	88	CBAL $E_{\text{cm}}^{\text{ee}}=10$ GeV	

$\Gamma(\mu^-\eta)/\Gamma_{\text{total}}$					$\Gamma_{191}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6.5\times10^{-8}</math></b>	90	MIYAZAKI	07	BELL 401 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.5\times10^{-7}$	90	AUBERT	07I	BABR 339 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.5\times10^{-7}$	90	ENARI	05	BELL 154 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.4\times10^{-7}$	90	ENARI	04	BELL 84.3 fb $^{-1}$ , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<9.6\times10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<7.3\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	

$\Gamma(e^-\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{192}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.8\times10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<4.6\times10^{-8}$	90	AUBERT	09w	BABR 451 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.3\times10^{-8}$	90	NISHIO	08	BELL 543 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.5\times10^{-7}$	90	YUSA	06	BELL 158 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<2.0\times10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<4.2\times10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98	
$<1.9\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<37\times10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^-\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{193}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.2\times10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<2.6\times10^{-8}$	90	AUBERT	09w	BABR 451 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.8\times10^{-8}$	90	NISHIO	08	BELL 543 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<2.0\times10^{-7}$	90	YUSA	06	BELL 158 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.3\times10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<5.7\times10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98	
$<2.9\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	
$<44\times10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV	

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(e^-\omega)/\Gamma_{\text{total}}$					$\Gamma_{194}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;4.8\times10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.1\times10^{-7}$	90	AUBERT	08K	BABR 384 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<1.8\times10^{-7}$	90	NISHIO	08	BELL 543 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	

$\Gamma(\mu^-\omega)/\Gamma_{\text{total}}$					$\Gamma_{195}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;4.7\times10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.0\times10^{-7}$	90	AUBERT	08K	BABR 384 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<8.9\times10^{-8}$	90	NISHIO	08	BELL 543 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	

$\Gamma(e^-K^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{196}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;3.2\times10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<5.9\times10^{-8}$	90	AUBERT	09w	BABR 451 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<7.8\times10^{-8}$	90	NISHIO	08	BELL 543 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<3.0\times10^{-7}$	90	YUSA	06	BELL 158 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<5.1\times10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$<6.3\times10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98	
$<3.8\times10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV	

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^- K^*(892)^0)/\Gamma_{\text{total}}$ $\Gamma_{197}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5.9 \times 10^{-8}</math></b>	90	NISHIO	08	BELL 543 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<7.2 \times 10^{-8}$	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.7 \times 10^{-7}$	90	AUBERT	09w	BABR 451 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<3.9 \times 10^{-7}$	90	YUSA	06	BELL 158 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<9.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<4.5 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10$ GeV

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$ $\Gamma_{198}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.4 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4.6 \times 10^{-8}$	90	AUBERT	09w	BABR 451 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.7 \times 10^{-8}$	90	NISHIO	08	BELL 543 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<4.0 \times 10^{-7}$	90	YUSA	06	BELL 158 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.1 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$ $\Gamma_{199}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;7.0 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<7.3 \times 10^{-8}$	90	AUBERT	09w	BABR 451 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.0 \times 10^{-7}$	90	NISHIO	08	BELL 543 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<4.0 \times 10^{-7}$	90	YUSA	06	BELL 158 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<8.7 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^- \eta'(958))/\Gamma_{\text{total}}$ $\Gamma_{200}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.6 \times 10^{-7}</math></b>	90	MIYAZAKI	07	BELL 401 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.4 \times 10^{-7}$	90	AUBERT	07i	BABR 339 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<10. \times 10^{-7}$	90	ENARI	05	BELL 154 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

### $\Gamma(\mu^- \eta'(958))/\Gamma_{\text{total}}$ $\Gamma_{201}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.3 \times 10^{-7}</math></b>	90	MIYAZAKI	07	BELL 401 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.4 \times 10^{-7}$	90	AUBERT	07i	BABR 339 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<4.7 \times 10^{-7}$	90	ENARI	05	BELL 154 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

### $\Gamma(e^- f_0(980) \rightarrow e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{202}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.2 \times 10^{-8}</math></b>	90	MIYAZAKI	09	BELL 671 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

### $\Gamma(\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{203}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.4 \times 10^{-8}</math></b>	90	MIYAZAKI	09	BELL 671 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

### $\Gamma(e^- \phi)/\Gamma_{\text{total}}$ $\Gamma_{204}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.1 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
<b><math>&lt;3.1 \times 10^{-8}</math></b>	90	AUBERT	09w	BABR 451 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<7.3 \times 10^{-8}$	90	NISHIO	08	BELL 543 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.3 \times 10^{-7}$	90	YUSA	06	BELL 158 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<6.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

### $\Gamma(\mu^- \phi)/\Gamma_{\text{total}}$ $\Gamma_{205}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;8.4 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL 854 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.9 \times 10^{-7}$	90	AUBERT	09w	BABR 451 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.3 \times 10^{-7}$	90	NISHIO	08	BELL 543 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.7 \times 10^{-7}$	90	YUSA	06	BELL 158 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<7.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

### $\Gamma(e^- e^+ e^-)/\Gamma_{\text{total}}$ $\Gamma_{206}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.7 \times 10^{-8}</math></b>	90	HAYASAKA	10	BELL 782 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.9 \times 10^{-8}$	90	LEES	10A	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<3.6 \times 10^{-8}$	90	MIYAZAKI	08	BELL 535 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<4.3 \times 10^{-8}$	90	AUBERT	07BK	BABR 376 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.0 \times 10^{-7}$	90	AUBERT	04J	BABR 91.5 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<3.5 \times 10^{-7}$	90	YUSA	04	BELL 87.1 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<0.33 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<1.3 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10$ GeV
$<2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4-10.9$
$<40 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8-6.8$ GeV

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^- \mu^+ \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{207}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.7 \times 10^{-8}</math></b>	90	HAYASAKA	10	BELL 782 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.2 \times 10^{-8}$	90	LEES	10A	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<4.1 \times 10^{-8}$	90	MIYAZAKI	08	BELL 535 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<3.7 \times 10^{-8}$	90	AUBERT	07BK	BABR 376 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<3.3 \times 10^{-7}$	90	AUBERT	04J	BABR 91.5 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.0 \times 10^{-7}$	90	YUSA	04	BELL 87.1 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<0.36 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<1.9 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10$ GeV
$<2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4-10.9$
$<33 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8-6.8$ GeV

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^+ \mu^- \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{208}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.7 \times 10^{-8}</math></b>	90	HAYASAKA	10	BELL 782 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.6 \times 10^{-8}$	90	LEES	10A	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.3 \times 10^{-8}$	90	MIYAZAKI	08	BELL 535 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<5.6 \times 10^{-8}$	90	AUBERT	07BK	BABR 376 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.3 \times 10^{-7}$	90	AUBERT	04J	BABR 91.5 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.0 \times 10^{-7}$	90	YUSA	04	BELL 87.1 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<0.35 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<1.8 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10$ GeV
$<1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4-10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^- e^+ e^-)/\Gamma_{\text{total}}$ $\Gamma_{209}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.8 \times 10^{-8}</math></b>	90	HAYASAKA	10	BELL 782 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.7 \times 10^{-8}$	90	LEES	10A	BABR 468 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.2 \times 10^{-8}$	90	MIYAZAKI	08	BELL 535 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<8.0 \times 10^{-8}$	90	AUBERT	07BK	BABR 376 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<2.7 \times 10^{-7}$	90	AUBERT	04J	BABR 91.5 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.9 \times 10^{-7}$	90	YUSA	04	BELL 87.1 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<1.7 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$<0.34 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<1.4 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10$ GeV
$<2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4-10.9$
$<44 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8-6.8$ GeV

<sup>1</sup> BARTELT 94 assume phase space decays.



See key on page 601

## Lepton Particle Listings

 $\tau$ 

$\Gamma(\mu^-\pi^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{222}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 8.6 \times 10^{-8}$	90	MIYAZAKI	13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 1.6 \times 10^{-7}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$< 2.7 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.6 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 7.5 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 8.7 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$< 11 \times 10^{-5}$	90	ALBRECHT	92k	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 7.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^-\pi^-K^+)/\Gamma_{\text{total}}$					$\Gamma_{223}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<4.5 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.0 \times 10^{-7}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13	
$<7.3 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$<3.2 \times 10^{-7}$	90	AUBERT,BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$<7.4 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$<1.5 \times 10^{-5}$	90	<sup>1</sup> BARTELT 94	CLEO	Repl. by BLISS 98	
$<7.7 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$	

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^+ \pi^- K^-)/\Gamma_{\text{total}}$					$\Gamma_{224}/\Gamma$
Test of lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<4.8 \times 10^{-8}$	90	MIYAZAKI	13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<9.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<2.9 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.0 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.0 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	ALBRECHT	92k	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<4.0 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- K_S^0 K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{225}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 8.0 \times 10^{-8}$	90	MIYAZAKI	10A	BELL	$671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 3.4 \times 10^{-6}$	90	CHEN	02c	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(\mu^- K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{226}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 4.4 \times 10^{-8}$	90	MIYAZAKI 13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 6.8 \times 10^{-8}$	90	MIYAZAKI 10	BELL	Repl. by MIYAZAKI 13	
$< 8.0 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 2.5 \times 10^{-7}$	90	AUBERT,BE 05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 15 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

$\Gamma(\mu^+ K^- K^-)/\Gamma_{\text{total}}$					$\Gamma_{227}/\Gamma$
Test of lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<4.7 \times 10^{-8}$	90	MIYAZAKI	13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<9.6 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<4.4 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<4.8 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$<6.0 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^- \pi^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{228}/\Gamma$	
Test of lepton family number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 6.5 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(\mu^- \pi^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{229}/\Gamma$	
Test of lepton family number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 14 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(e^- \eta \eta)/\Gamma_{\text{total}}$					$\Gamma_{230}/\Gamma$	
Test of lepton family number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 35 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(\mu^- \eta \eta)/\Gamma_{\text{total}}$					$\Gamma_{231}/\Gamma$	
Test of lepton family number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 60 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(e^- \pi^0 \eta)/\Gamma_{\text{total}}$					$\Gamma_{232}/\Gamma$	
Test of lepton family number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 24 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(\mu^- \pi^0 \eta)/\Gamma_{\text{total}}$					$\Gamma_{233}/\Gamma$	
Test of lepton family number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 22 \times 10^{-6}$	90	BONVICINI 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(p \mu^- \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{234}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 4.4 \times 10^{-7}$	90	AAIJ 13AH	LHCB	$1.0 \text{ fb}^{-1}, \sqrt{s} = 7 \text{ TeV}$		

$\Gamma(\bar{p} \mu^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{235}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 3.3 \times 10^{-7}$	90	AAIJ 13AH	LHCB	$1.0 \text{ fb}^{-1}, \sqrt{s} = 7 \text{ TeV}$		

$\Gamma(\bar{p}\gamma)/\Gamma_{\text{total}}$					$\Gamma_{236}/\Gamma$
Test of lepton number and baryon number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 3.5 \times 10^{-6}$	90	GODANG	99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$< 29 \times 10^{-5}$	90	ALBRECHT	92k	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

$\Gamma(\bar{p}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{237}/\Gamma$
Test of lepton number and baryon number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<15 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<66 \times 10^{-5}$	90	ALBRECHT 92k	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	

$\Gamma(\bar{p} 2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{238}/\Gamma$	
Test of lepton number and baryon number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 33 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(\overline{p}\eta)/\Gamma_{\text{total}}$					$\Gamma_{239}/\Gamma$
Test of lepton number and baryon number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 8.9 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 130 \times 10^{-5}$	90	ALBRECHT 92k	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	

$\Gamma(\bar{p} \pi^0 \eta)/\Gamma_{\text{total}}$					$\Gamma_{240}/\Gamma$	
Test of lepton number and baryon number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 27 \times 10^{-6}$	90	GODANG 99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(\Lambda \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{241}/\Gamma$	
Test of lepton number and baryon number conservation.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 0.72 \times 10^{-7}$	90	MIYAZAKI 06	BELL	$154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		

$\Gamma(\bar{\Lambda} \pi^-)/\Gamma_{\text{total}}$
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# Lepton Particle Listings

$\tau$

$\Gamma(e^- \text{light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$					$\Gamma_{243}/\Gamma_5$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.015	95	<sup>1</sup> ALBRECHT	95G ARG	$E_{\text{cm}}^{\text{e}} = 9.4\text{--}10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.018	95	<sup>2</sup> ALBRECHT	90E ARG	$E_{\text{cm}}^{\text{e}} = 9.4\text{--}10.6$ GeV	
<0.040	95	<sup>3</sup> BALTRUSAITIS	85 MRK3	$E_{\text{cm}}^{\text{e}} = 3.77$ GeV	
<sup>1</sup> ALBRECHT 95G limit holds for bosons with mass < 0.4 GeV. The limit rises to 0.036 for a mass of 1.0 GeV, then falls to 0.006 at the upper mass limit of 1.6 GeV.					
<sup>2</sup> ALBRECHT 90E limit applies for spinless boson with mass < 100 MeV, and rises to 0.050 for mass = 500 MeV.					
<sup>3</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass < 100 MeV.					

$\Gamma(\mu^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$					$\Gamma_{244}/\Gamma_5$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.026</b>	95	<sup>1</sup> ALBRECHT	95G ARG	$E_{\text{cm}}^e = 9.4\text{--}10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.033	95	<sup>2</sup> ALBRECHT	90E ARG	$E_{\text{cm}}^e = 9.4\text{--}10.6$ GeV	
<0.125	95	<sup>3</sup> BALTRUSAITIS	85 MRK3	$E_{\text{cm}}^e = 3.77$ GeV	
<sup>1</sup> ALBRECHT 95G limit holds for bosons with mass < 1.3 GeV. The limit rises to 0.034 for a mass of 1.4 GeV, then falls to 0.003 at the upper mass limit of 1.6 GeV.					
<sup>2</sup> ALBRECHT 90E limit applies for spinless boson with mass < 100 MeV, and rises to 0.071 for mass = 500 MeV.					
<sup>3</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass < 100 MeV.					

## $\tau$ -DECAY PARAMETERS $\tau$ -LEPTON DECAY PARAMETERS

Updated August 2011 by A. Stahl (RWTH Aachen).

The purpose of the measurements of the decay parameters (also known as Michel parameters) of the  $\tau$  is to determine the structure (spin and chirality) of the current mediating its decays.

**Leptonic Decays:** The Michel parameters are extracted from the energy spectrum of the charged daughter lepton  $\ell = e, \mu$  in the decays  $\tau \rightarrow \ell \nu_\ell \nu_\tau$ . Ignoring radiative corrections, neglecting terms of order  $(m_\ell/m_\tau)^2$  and  $(m_\tau/\sqrt{s})^2$ , and setting the neutrino masses to zero, the spectrum in the laboratory frame reads

$$\frac{d\Gamma}{dx} = \frac{G_{\tau\ell}^2 m_\tau^5}{192 \pi^3} \times \left\{ f_0(x) + \rho f_1(x) + \eta \frac{m_\ell}{m_\tau} f_2(x) - P_\tau [\xi g_1(x) + \xi \delta g_2(x)] \right\}, \quad (1)$$

with

$$\begin{aligned} f_0(x) &= 2 - 6x^2 + 4x^3 \\ f_1(x) &= -\frac{4}{9} + 4x^2 - \frac{32}{9}x^3 & g_1(x) &= -\frac{2}{3} + 4x - 6x^2 + \frac{8}{3}x^3 \\ f_2(x) &= 12(1-x)^2 & g_2(x) &= \frac{4}{9} - \frac{16}{3}x + 12x^2 - \frac{64}{9}x^3. \end{aligned}$$

The quantity  $x$  is the fractional energy of the daughter lepton  $\ell$ , i.e.,  $x = E_\ell/E_{\ell, \text{max}} \approx E_\ell/(\sqrt{s}/2)$  and  $P_\tau$  is the polarization of the tau leptons. The integrated decay width is given by

$$\Gamma = \frac{G_{\tau\ell}^2 m_\tau^5}{192 \pi^3} \left( 1 + 4\eta \frac{m_\ell}{m_\tau} \right). \quad (2)$$

The situation is similar to muon decays  $\mu \rightarrow e \nu_e \nu_\mu$ . The generalized matrix element with the couplings  $g_{\tau\ell}^\gamma$  and their relations to the Michel parameters  $\rho$ ,  $\eta$ ,  $\xi$ , and  $\delta$  have been described in the “Note on Muon Decay Parameters.” The Standard Model expectations are 3/4, 0, 1, and 3/4, respectively. For more details, see Ref. 1.

**Hadronic Decays:** In the case of hadronic decays  $\tau \rightarrow h \nu_\tau$ , with  $h = \pi, \rho$ , or  $a_1$ , the ansatz is restricted to purely vectorial currents. The matrix element is

$$\frac{G_{\tau h}}{\sqrt{2}} \sum_{\lambda=R,L} g_\lambda \langle \bar{\Psi}_\omega(\nu_\tau) | \gamma^\mu | \Psi_\lambda(\tau) \rangle J_\mu^h \quad (3)$$

with the hadronic current  $J_\mu^h$ . The neutrino chirality  $\omega$  is uniquely determined from  $\lambda$ . The spectrum depends only on a single parameter  $\xi_h$

$$\frac{d^n \Gamma}{dx_1 dx_2 \dots dx_n} = f(\vec{x}) + \xi_h P_\tau g(\vec{x}), \quad (4)$$

with  $f$  and  $g$  being channel-dependent functions of the  $n$  observables  $\vec{x} = (x_1, x_2, \dots, x_n)$  (see Ref. 2). The parameter  $\xi_h$  is related to the couplings through

$$\xi_h = |g_L|^2 - |g_R|^2. \quad (5)$$

$\xi_h$  is the negative of the chirality of the  $\tau$  neutrino in these decays. In the Standard Model,  $\xi_h = 1$ . Also included in the Data Listings for  $\xi_h$  are measurements of the neutrino helicity which coincide with  $\xi_h$ , if the neutrino is massless (ASNER 00, ACKERSTAFF 97R, AKERS 95P, ALBRECHT 93C, and ALBRECHT 90I).

**Combination of Measurements:** The individual measurements are combined, taking into account the correlations between the parameters. In a first fit, universality between the two leptonic decays, and between all hadronic decays, is assumed. A second fit is made without these assumptions. The results of the two fits are provided as OUR FIT in the Data Listings below in the tables whose title includes “(e or mu)” or “(all hadronic modes),” and “(e),” “(mu)” etc., respectively. The measurements show good agreement with the Standard Model. The  $\chi^2$  values with respect to the Standard model predictions are 24.1 for 41 degrees of freedom and 26.8 for 56 degrees of freedom, respectively. The correlations are reduced through this combination to less than 20%, with the exception of  $\rho$  and  $\eta$  which are correlated by +23%, for the fit with universality and by +70% for  $\tau \rightarrow \mu \nu_\mu \nu_\tau$ .

**Model-independent Analysis:** From the Michel parameters, limits can be derived on the couplings  $g_{e\lambda}^\kappa$  without further model assumptions. In the Standard model  $g_{LL}^V = 1$  (leptonic decays), and  $g_L = 1$  (hadronic decays) and all other couplings vanish. First, the partial decay widths have to be compared to the Standard Model predictions to derive limits on the normalization of the couplings  $A_x = G_{\tau x}^2/G_F^2$  with Fermi’s constant  $G_F$ :

$$A_e = 1.0029 \pm 0.0046,$$

$$A_\mu = 0.981 \pm 0.018,$$

$$A_\pi = 1.0020 \pm 0.0073. \quad (6)$$

Then limits on the couplings (95% CL) can be extracted (see Ref. 3 and Ref. 4). Without the assumption of universality, the limits given in Table 1 are derived.



**Table 1:** Coupling constants  $g_{e\mu}^\gamma$ . 95% confidence level experimental limits. The limits include the quoted values of  $A_e$ ,  $A_\mu$ , and  $A_\pi$  and assume  $A_\rho = A_{a_1} = 1$ .

$\tau \rightarrow e\nu_e\nu_\tau$		
$ g_{RR}^S  < 0.70$	$ g_{RR}^V  < 0.17$	$ g_{RR}^T  \equiv 0$
$ g_{LR}^S  < 0.99$	$ g_{LR}^V  < 0.13$	$ g_{LR}^T  < 0.082$
$ g_{RL}^S  < 2.01$	$ g_{RL}^V  < 0.52$	$ g_{RL}^T  < 0.51$
$ g_{LL}^S  < 2.01$	$ g_{LL}^V  < 1.005$	$ g_{LL}^T  \equiv 0$
$\tau \rightarrow \mu\nu_\mu\nu_\tau$		
$ g_{RR}^S  < 0.72$	$ g_{RR}^V  < 0.18$	$ g_{RR}^T  \equiv 0$
$ g_{LR}^S  < 0.95$	$ g_{LR}^V  < 0.12$	$ g_{LR}^T  < 0.079$
$ g_{RL}^S  < 2.01$	$ g_{RL}^V  < 0.52$	$ g_{RL}^T  < 0.51$
$ g_{LL}^S  < 2.01$	$ g_{LL}^V  < 1.005$	$ g_{LL}^T  \equiv 0$
$\tau \rightarrow \pi\nu_\tau$		
$ g_R^V  < 0.15$	$ g_L^V  > 0.992$	
$\tau \rightarrow \rho\nu_\tau$		
$ g_R^V  < 0.10$	$ g_L^V  > 0.995$	
$\tau \rightarrow a_1\nu_\tau$		
$ g_R^V  < 0.16$	$ g_L^V  > 0.987$	

**Model-dependent Interpretation:** More stringent limits can be derived assuming specific models. For example, in the framework of a two Higgs doublet model, the measurements correspond to a limit of  $m_{H^\pm} > 1.9 \text{ GeV} \times \tan\beta$  on the mass of the charged Higgs boson, or a limit of 253 GeV on the mass of the second  $W$  boson in left-right symmetric models for arbitrary mixing (both 95% CL). See Ref. 4 and Ref. 5.

## Footnotes and References

1. F. Scheck, Phys. Reports **44**, 187 (1978);  
W. Fetscher and H.J. Gerber in *Precision Tests of the Standard Model*, edited by P. Langacker, World Scientific, 1993;  
A. Stahl, *Physics with  $\tau$  Leptons*, Springer Tracts in Modern Physics.
2. M. Davier *et al.*, Phys. Lett. **B306**, 411 (1993).
3. OPAL Collab., K. Ackerstaff *et al.*, Eur. Phys. J. **C8**, 3 (1999).
4. A. Stahl, Nucl. Phys. (Proc. Supp.) **B76**, 173 (1999).
5. M.-T. Dova *et al.*, Phys. Rev. **D58**, 015005 (1998);  
T. Hebbeker and W. Lohmann, Z. Phys. **C74**, 399 (1997);  
A. Pich and J.P. Silva, Phys. Rev. **D52**, 4006 (1995).

## $\rho(e \text{ or } \mu)$ PARAMETER

(V–A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.745 ± 0.008 OUR FIT</b>				
<b>0.749 ± 0.008 OUR AVERAGE</b>				
0.742 ± 0.014 ± 0.006	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.775 ± 0.023 ± 0.020	36k	ABREU	00L DLPH	1992–1995 runs
0.781 ± 0.028 ± 0.018	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.762 ± 0.035	54k	ACCIARRI	98R L3	1991–1995 LEP runs

0.731 ± 0.031		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.72 ± 0.09 ± 0.03		<sup>2</sup> ABE	97o SLD	1993–1995 SLC runs
0.747 ± 0.010 ± 0.006	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.79 ± 0.10 ± 0.10	373k	FORD	87B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
0.71 ± 0.09 ± 0.03	142k	BEHREND	85 CLEO	$e^+e^-$ near $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.735 ± 0.013 ± 0.008	31k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.794 ± 0.039 ± 0.031	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
0.732 ± 0.034 ± 0.020	8.2k	<sup>3</sup> ALBRECHT	95 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.738 ± 0.038		<sup>4</sup> ALBRECHT	95c ARG	Repl. by ALBRECHT 98
0.751 ± 0.039 ± 0.022		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.742 ± 0.035 ± 0.020	8000	ALBRECHT	90E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95c, ALBRECHT 93g, and ALBRECHT 94e. ALBRECHT 98 use tau pair events of the type  $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\nu_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 97o assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\rho$  value of  $0.69 \pm 0.13 \pm 0.05$ .

<sup>3</sup> Value is from a simultaneous fit for the  $\rho$  and  $\eta$  decay parameters to the lepton energy spectrum. Not independent of ALBRECHT 90E ( $\rho(e \text{ or } \mu)$ ) value which assumes  $\eta = 0$ . Result is strongly correlated with ALBRECHT 95c.

<sup>4</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95c, ALBRECHT 93g, and ALBRECHT 94e.

## $\rho(e)$ PARAMETER

(V–A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.747 ± 0.010 OUR FIT</b>				
<b>0.744 ± 0.010 OUR AVERAGE</b>				
0.747 ± 0.019 ± 0.014	44k	HEISTER	01E ALEP	1991–1995 LEP runs
0.744 ± 0.036 ± 0.037	17k	ABREU	00L DLPH	1992–1995 runs
0.779 ± 0.047 ± 0.029	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.68 ± 0.04 ± 0.07		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.71 ± 0.14 ± 0.05		ABE	97o SLD	1993–1995 SLC runs
0.747 ± 0.012 ± 0.004	34k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.735 ± 0.036 ± 0.020	4.7k	<sup>2</sup> ALBRECHT	95 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.79 ± 0.08 ± 0.06	3230	<sup>3</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
0.64 ± 0.06 ± 0.07	2753	JANSSEN	89 CBAL	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
0.62 ± 0.17 ± 0.14	1823	FORD	87B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
0.60 ± 0.13	699	BEHREND	85 CLEO	$e^+e^-$ near $\Upsilon(4S)$
0.72 ± 0.10 ± 0.11	594	BACINO	79B DLCO	$E_{\text{cm}}^{\text{ee}} = 3.5\text{--}7.4 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.732 ± 0.014 ± 0.009	19k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.793 ± 0.050 ± 0.025		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.747 ± 0.045 ± 0.028	5106	ALBRECHT	90E ARG	Repl. by ALBRECHT 95
<sup>1</sup> ALBRECHT 98 use tau pair events of the type $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\nu_\tau)$ , and their charged conjugates.				
<sup>2</sup> ALBRECHT 95 use tau pair events of the type $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(h^+h^-\pi^0\nu_\tau)$ and their charged conjugates.				
<sup>3</sup> ALBRECHT 93G use tau pair events of the type $\tau^-\tau^+ \rightarrow (\mu^-\bar{\nu}_\mu\nu_\tau)(e^+\nu_e\nu_\tau)$ and their charged conjugates.				

## $\rho(\mu)$ PARAMETER

(V–A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.763 ± 0.020 OUR FIT</b>				
<b>0.770 ± 0.022 OUR AVERAGE</b>				
0.776 ± 0.045 ± 0.019	46k	HEISTER	01E ALEP	1991–1995 LEP runs
0.999 ± 0.098 ± 0.045	22k	ABREU	00L DLPH	1992–1995 runs
0.777 ± 0.044 ± 0.016	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.69 ± 0.06 ± 0.06		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.54 ± 0.28 ± 0.14		ABE	97o SLD	1993–1995 SLC runs
0.750 ± 0.017 ± 0.045	22k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.76 ± 0.07 ± 0.08	3230	ALBRECHT	93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
0.734 ± 0.055 ± 0.027	3041	ALBRECHT	90E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
0.89 ± 0.14 ± 0.08	1909	FORD	87B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
0.81 ± 0.13	727	BEHREND	85 CLEO	$e^+e^-$ near $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.747 ± 0.048 ± 0.044	13k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.693 ± 0.057 ± 0.028		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
<sup>1</sup> ALBRECHT 98 use tau pair events of the type $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\nu_\tau)$ , and their charged conjugates.				

## $\xi(e \text{ or } \mu)$ PARAMETER

(V–A) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.985 ± 0.030 OUR FIT</b>				
<b>0.981 ± 0.031 OUR AVERAGE</b>				
0.986 ± 0.068 ± 0.031	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.929 ± 0.070 ± 0.030	36k	ABREU	00L DLPH	1992–1995 runs
0.98 ± 0.22 ± 0.10	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.70 ± 0.16	54k	ACCIARRI	98R L3	1991–1995 LEP runs
1.03 ± 0.11		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
1.05 ± 0.35 ± 0.04		<sup>2</sup> ABE	97o SLD	1993–1995 SLC runs
1.007 ± 0.040 ± 0.015	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

# Lepton Particle Listings

## $\tau$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.94 \pm 0.21 \pm 0.07$	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
$0.97 \pm 0.14$		<sup>3</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
$1.18 \pm 0.15 \pm 0.16$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
$0.90 \pm 0.15 \pm 0.10$	3230	<sup>4</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \overline{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 97o assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\xi$  value of  $1.02 \pm 0.36 \pm 0.05$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \overline{\nu}_\tau)$  and their charged conjugates.

<sup>4</sup> ALBRECHT 93G measurement determines  $|\xi|$  for the case  $\xi(e) = \xi(\mu)$ , but the authors point out that other LEP experiments determine the sign to be positive.

## $\xi(e)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.994 \pm 0.040</math> OUR FIT</b>				
<b><math>1.00 \pm 0.04</math> OUR AVERAGE</b>				
$1.011 \pm 0.094 \pm 0.038$	44k	HEISTER	01E ALEP	1991–1995 LEP runs
$1.01 \pm 0.12 \pm 0.05$	17k	ABREU	00L DLPH	1992–1995 runs
$1.13 \pm 0.39 \pm 0.14$	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$1.11 \pm 0.20 \pm 0.08$		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6$ GeV
$1.16 \pm 0.52 \pm 0.06$		ABE	97o SLD	1993–1995 SLC runs
$0.979 \pm 0.048 \pm 0.016$	34k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.03 \pm 0.23 \pm 0.09$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \overline{\nu}_\tau)$ , and their charged conjugates.

## $\xi(\mu)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.030 \pm 0.059</math> OUR FIT</b>				
<b><math>1.06 \pm 0.06</math> OUR AVERAGE</b>				
$1.030 \pm 0.120 \pm 0.050$	46k	HEISTER	01E ALEP	1991–1995 LEP runs
$1.16 \pm 0.19 \pm 0.06$	22k	ABREU	00L DLPH	1992–1995 runs
$0.79 \pm 0.41 \pm 0.09$	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$1.26 \pm 0.27 \pm 0.14$		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6$ GeV
$0.75 \pm 0.50 \pm 0.14$		ABE	97o SLD	1993–1995 SLC runs
$1.054 \pm 0.069 \pm 0.047$	22k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.23 \pm 0.22 \pm 0.10$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \overline{\nu}_\tau)$ , and their charged conjugates.

## $\eta(e \text{ or } \mu)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.013 \pm 0.020</math> OUR FIT</b>				
<b><math>0.015 \pm 0.021</math> OUR AVERAGE</b>				
$0.012 \pm 0.026 \pm 0.004$	81k	HEISTER	01E ALEP	1991–1995 LEP runs
$-0.005 \pm 0.036 \pm 0.037$		ABREU	00L DLPH	1992–1995 runs
$0.027 \pm 0.055 \pm 0.005$	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$0.27 \pm 0.14$	54k	ACCIARRI	98R L3	1991–1995 LEP runs
$-0.13 \pm 0.47 \pm 0.15$		ABE	97o SLD	1993–1995 SLC runs
$-0.015 \pm 0.061 \pm 0.062$	31k	AMMAR	97B CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$0.03 \pm 0.18 \pm 0.12$	8.2k	ALBRECHT	95 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.25 \pm 0.17 \pm 0.11$	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
$-0.04 \pm 0.15 \pm 0.11$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

## $\eta(\mu)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.094 \pm 0.073</math> OUR FIT</b>				
<b><math>0.17 \pm 0.15</math> OUR AVERAGE</b>				Error includes scale factor of 1.2.
$0.160 \pm 0.150 \pm 0.060$	46k	HEISTER	01E ALEP	1991–1995 LEP runs
$0.72 \pm 0.32 \pm 0.15$		ABREU	00L DLPH	1992–1995 runs
$-0.59 \pm 0.82 \pm 0.45$		<sup>1</sup> ABE	97o SLD	1993–1995 SLC runs
$0.010 \pm 0.149 \pm 0.171$	13k	<sup>2</sup> AMMAR	97B CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.010 \pm 0.065 \pm 0.001$	27k	<sup>3</sup> ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$-0.24 \pm 0.23 \pm 0.18$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

<sup>1</sup> Highly correlated (corr. = 0.92) with ABE 97o  $\rho(\mu)$  measurement.

<sup>2</sup> Highly correlated (corr. = 0.949) with AMMAR 97B  $\rho(\mu)$  value.

<sup>3</sup> ACKERSTAFF 99D result is dominated by a constraint on  $\eta$  from the OPAL measurements of the  $\tau$  lifetime and  $B(\tau^- \rightarrow \mu^- \overline{\nu}_\mu \nu_\tau)$  assuming lepton universality for the total coupling strength.

## $(\delta\xi)(e \text{ or } \mu)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.746 \pm 0.021</math> OUR FIT</b>				
<b><math>0.744 \pm 0.022</math> OUR AVERAGE</b>				
$0.776 \pm 0.045 \pm 0.024$	81k	HEISTER	01E ALEP	1991–1995 LEP runs
$0.779 \pm 0.070 \pm 0.028$	36k	ABREU	00L DLPH	1992–1995 runs
$0.65 \pm 0.14 \pm 0.07$	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$0.70 \pm 0.11$	54k	ACCIARRI	98R L3	1991–1995 LEP runs
$0.63 \pm 0.09$		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6$ GeV
$0.88 \pm 0.27 \pm 0.04$		<sup>2</sup> ABE	97o SLD	1993–1995 SLC runs
$0.745 \pm 0.026 \pm 0.009$	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.81 \pm 0.14 \pm 0.06$	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
$0.65 \pm 0.12$		<sup>3</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
$0.88 \pm 0.11 \pm 0.07$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \overline{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 97o assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $(\delta\xi)$  value of  $0.87 \pm 0.27 \pm 0.04$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \overline{\nu}_\tau)$  and their charged conjugates.

## $(\delta\xi)(e)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.734 \pm 0.028</math> OUR FIT</b>				
<b><math>0.731 \pm 0.029</math> OUR AVERAGE</b>				
$0.778 \pm 0.066 \pm 0.024$	44k	HEISTER	01E ALEP	1991–1995 LEP runs
$0.85 \pm 0.12 \pm 0.04$	17k	ABREU	00L DLPH	1992–1995 runs
$0.72 \pm 0.31 \pm 0.14$	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$0.56 \pm 0.14 \pm 0.06$		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6$ GeV
$0.85 \pm 0.43 \pm 0.08$		ABE	97o SLD	1993–1995 SLC runs
$0.720 \pm 0.032 \pm 0.010$	34k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.11 \pm 0.17 \pm 0.07$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \overline{\nu}_\tau)$ , and their charged conjugates.

## $(\delta\xi)(\mu)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.778 \pm 0.037</math> OUR FIT</b>				
<b><math>0.79 \pm 0.04</math> OUR AVERAGE</b>				
$0.786 \pm 0.066 \pm 0.028$	46k	HEISTER	01E ALEP	1991–1995 LEP runs
$0.86 \pm 0.13 \pm 0.04$	22k	ABREU	00L DLPH	1992–1995 runs
$0.63 \pm 0.23 \pm 0.05$	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
$0.73 \pm 0.18 \pm 0.10$		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6$ GeV
$0.82 \pm 0.32 \pm 0.07$		ABE	97o SLD	1993–1995 SLC runs
$0.786 \pm 0.041 \pm 0.032$	22k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.71 \pm 0.14 \pm 0.06$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \overline{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \overline{\nu}_\tau)$ , and their charged conjugates.

## $\xi(\pi)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $\xi(\pi) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.993 \pm 0.022</math> OUR FIT</b>				
<b><math>0.994 \pm 0.023</math> OUR AVERAGE</b>				
$0.994 \pm 0.020 \pm 0.014$	27k	HEISTER	01E ALEP	1991–1995 LEP runs
$0.81 \pm 0.17 \pm 0.02$		ABE	97o SLD	1993–1995 SLC runs
$1.03 \pm 0.06 \pm 0.04$	2.0k	COAN	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.987 \pm 0.057 \pm 0.027$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
$0.95 \pm 0.11 \pm 0.05$		<sup>1</sup> BUSKULIC	94D ALEP	1990+1991 LEP run

<sup>1</sup> Superseded by BUSKULIC 95D.

## $\xi(\rho)$ PARAMETER

( $V\text{--}A$ ) theory predicts  $\xi(\rho) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.994 \pm 0.008</math> OUR FIT</b>				
<b><math>0.994 \pm 0.009</math> OUR AVERAGE</b>				
$0.987 \pm 0.012 \pm 0.011$	59k	HEISTER	01E ALEP	1991–1995 LEP runs
$0.99 \pm 0.12 \pm 0.04$		ABE	97o SLD	1993–1995 SLC runs
$0.995 \pm 0.010 \pm 0.003$	66k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
$1.022 \pm 0.028 \pm 0.030$	1.7k	<sup>1</sup> ALBRECHT	94E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.045 \pm 0.058 \pm 0.032$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
$1.03 \pm 0.11 \pm 0.05$		<sup>2</sup> BUSKULIC	94D ALEP	1990+1991 LEP run

<sup>1</sup> ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result.

<sup>2</sup> Superseded by BUSKULIC 95D.





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# Lepton Particle Listings

## Heavy Charged Lepton Searches, Neutrino Properties

<sup>7</sup> Assumes associated neutrino is approximately massless.  
<sup>8</sup> ADEVA 85 analyze one-isolated-muon data and sensitive to  $\tau < 10$  nanosec. Assume  $B(\text{lepton}) = 0.30$ .  $E_{\text{cm}} = 40\text{--}47$  GeV.  
<sup>9</sup> BARTEL 83 limit is from PETRA  $e^+e^-$  experiment with average  $E_{\text{cm}} = 34.2$  GeV.  
<sup>10</sup> BERGER 81b is DESY DORIS and PETRA experiment. Looking for  $e^+e^- \rightarrow L^+L^-$ .  
<sup>11</sup> BRANDELIK 81 is DESY-PETRA experiment. Looking for  $e^+e^- \rightarrow L^+L^-$ .  
<sup>12</sup> AZIMOV 80 estimated probabilities for  $M+N$  type events in  $e^+e^- \rightarrow L^+L^-$  deducing semi-hadronic decay multiplicities of  $L$  from  $e^+e^-$  annihilation data at  $E_{\text{cm}} = (2/3)m_L$ .  
Obtained above limit comparing these with  $e^+e^-$  data (BRANDELIK 80).  
<sup>13</sup> BARBER 80b looked for  $e^+e^- \rightarrow L^+L^-$ ,  $L \rightarrow \nu_L^+ X$  with MARK-J at DESY-PETRA.  
<sup>14</sup> ROTHE 69 examines previous data on  $\mu$  pair production and  $\pi$  and  $K$  decays.

### Stable Charged Heavy Lepton ( $L^\pm$ ) MASS LIMITS

VALUE (GeV)	CL%	DOCUMENT ID	TECN
>102.6	95	ACHARD 01B L3	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
> 28.2	95	<sup>15</sup> ADACHI 90C TOPZ	
none 18.5–42.8	95	AKRAWY 90a OPAL	
> 26.5	95	DECAMP 90F ALEP	
none $m_{\mu^-} \rightarrow 36.3$	95	SODERSTROM90 MRK2	

<sup>15</sup> ADACHI 90C put lower limits on the mass of stable charged particles with electric charge  $Q$  satisfying  $2/3 < Q/e < 4/3$  and with spin 0 or 1/2. We list here the special case for a stable charged heavy lepton.

### Charged Long-Lived Heavy Lepton MASS LIMITS

VALUE (GeV)	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>574	95	CHATRCHYAN13AB CMS			Leptons singlet model
>102.0	95	ABBIENDI 03L OPAL			pair produced in $e^+e^-$
> 0.1	16	ANSORGE 73B HBC	—		Long-lived
none 0.55–4.5	17	BUSHNIN 73 CNTR	—		Long-lived
none 0.2–0.92	18	BARNA 68 CNTR	—		Long-lived
none 0.97–1.03	18	BARNA 68 CNTR	—		Long-lived

<sup>16</sup> ANSORGE 73B looks for electron pair production and electron-like Bremsstrahlung.  
<sup>17</sup> BUSHNIN 73 is SERPUKHOV 70 GeV  $p$  experiment. Masses assume mean life above  $7 \times 10^{-10}$  and  $3 \times 10^{-8}$  respectively. Calculated from cross section (see “Charged Quasi-Stable Lepton Production Differential Cross Section” below) and 30 GeV muon pair production data.  
<sup>18</sup> BARNA 68 is SLAC photoproduction experiment.

### Doubly-Charged Heavy Lepton MASS LIMITS

VALUE (GeV)	CL%	DOCUMENT ID	TECN	CHG
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 1–9 GeV	90	<sup>19</sup> CLARK 81 SPEC ++		
<sup>19</sup> CLARK 81 is FNAL experiment with 209 GeV muons. Bounds apply to $\mu\mu$ which couples with full weak strength to muon. See also section on “Doubly-Charged Lepton Production Cross Section.”				

### Doubly-Charged Lepton Production Cross Section ( $\mu N$ Scattering)

VALUE (cm <sup>2</sup> )	EVTS	DOCUMENT ID	TECN	CHG
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<6. $\times 10^{-38}$	0	<sup>20</sup> CLARK 81 SPEC ++		
<sup>20</sup> CLARK 81 is FNAL experiment with 209 GeV muon. Looked for $\mu^+ \text{nucleon} \rightarrow \bar{\mu}_P^0 X$ , $\bar{\mu}_P^0 \rightarrow \mu^+ \mu^- \bar{\nu}_\mu$ and $\mu^+ n \rightarrow \mu_P^{++} X, \mu_P^{++} \rightarrow 2\mu^+ \nu_\mu$ . Above limits are for $\sigma \times \text{BR}$ taken from their mass-dependence plot figure 2.				

### REFERENCES FOR Heavy Charged Lepton Searches

CHATRCHYAN 13AB JHEP 1307 122	S. Chatrchyan <i>et al.</i> (CMS Collab.)
ABBIENDI 03L PL B572 8	G. Abbiendi <i>et al.</i> (OPAL Collab.)
ACHARD 01B PL B517 75	P. Achard <i>et al.</i> (L3 Collab.)
ACKERSTAFF 98C EPJ C1 45	K. Ackerstaff <i>et al.</i> (OPAL Collab.)
ACCIARRI 96G PL B377 304	M. Acciarri <i>et al.</i> (L3 Collab.)
ALEXANDER 96P PL B385 433	G. Alexander <i>et al.</i> (OPAL Collab.)
BUSKULIC 96S PL B384 439	D. Buskulić <i>et al.</i> (ALEPH Collab.)
AHMED 94 PL B340 205	T. Ahmed <i>et al.</i> (H1 Collab.)
KIM 91B UMP A6 2583	G.N. Kim <i>et al.</i> (AMY Collab.)
ADACHI 90C PL B244 352	I. Adachi <i>et al.</i> (TOPAZ Collab.)
AKRAWY 90G PL B240 250	M.Z. Akrawy <i>et al.</i> (OPAL Collab.)
AKRAWY 90O PL B252 290	M.Z. Akrawy <i>et al.</i> (OPAL Collab.)
DECAMP 90F PL B236 511	D. Decamp <i>et al.</i> (ALEPH Collab.)
RILES 90 PR D42 1	K. Riles <i>et al.</i> (Mark II Collab.)
SODERSTROM 90 PRL 64 2980	E. Soderstrom <i>et al.</i> (Mark II Collab.)
STOKER 89 PR D39 1811	D.P. Stoker <i>et al.</i> (Mark II Collab.)
ABE 88 PRL 61 915	K. Abe <i>et al.</i> (VENUS Collab.)
ADACHI 88B PR D37 1339	I. Adachi <i>et al.</i> (TOPAZ Collab.)
BEHREND 88C ZPHY C41 7	H.J. Behrend <i>et al.</i> (CELLO Collab.)
ALBAJAR 87B PL B185 241	C. Albajar <i>et al.</i> (UA1 Collab.)
ADEVA 85 PL 152B 439	B. Adeva <i>et al.</i> (Mark-J Collab.)
Also PRPL 109 131	B. Adeva <i>et al.</i> (Mark-J Collab.)
BARTEL 83 PL 123B 353	W. Bartel <i>et al.</i> (JADE Collab.)
BERGER 81B PL 99B 489	C. Berger <i>et al.</i> (PLUTO Collab.)
BRANDELIK 81 PL 99B 163	R. Brandelik <i>et al.</i> (TASSO Collab.)
CLARK 81 PRL 46 2799	A.R. Clark <i>et al.</i> (Mark II Collab.)
Also PR D25 2762	W.H. Smith <i>et al.</i> (UCB, LBL, FNAL+)
AZIMOV 80 JETPL 32 664	Y.I. Azimov, V.A. Khoze (Mark II Collab.)
Translated from ZETFP 32 677	
BARBER 80B PRL 45 1904	D.P. Barber <i>et al.</i> (Mark-J Collab.)
BRANDELIK 80 PL 92B 199	R. Brandelik <i>et al.</i> (TASSO Collab.)
ANSORGE 73B PR D7 26	R.E. Ansorge <i>et al.</i> (CAVE Collab.)
BUSHNIN 73 NP B58 476	Y.B. Bushnin <i>et al.</i> (SERP Collab.)
Also PL 42B 136	S.V. Golovkin <i>et al.</i> (SERP Collab.)
ROTHE 69 NP B10 241	K.W. Rothe, A.M. Wolsky (PERN Collab.)
BARNA 68 PR 173 1391	A. Barna <i>et al.</i> (SLAC, STAN)

### OTHER RELATED PAPERS

PERL 81 SLAC-PUB-2752 M.L. Perl (SLAC)  
Physics in Collision Conference.

## Neutrino Properties

### INTRODUCTION TO THE NEUTRINO PROPERTIES LISTINGS

Revised August 2013 by P. Vogel (Caltech) and A. Piepke (University of Alabama).

The following Listings concern measurements of various properties of neutrinos. Nearly all of the measurements, all of which so far are limits, actually concern superpositions of the mass eigenstates  $\nu_i$ , which are in turn related to the weak eigenstates  $\nu_\ell$ , via the neutrino mixing matrix

$$|\nu_\ell\rangle = \sum_i U_{\ell i} |\nu_i\rangle.$$

In the analogous case of quark mixing via the CKM matrix, the smallness of the off-diagonal terms (small mixing angles) permits a “dominant eigenstate” approximation. However, the results of neutrino oscillation searches show that the mixing matrix contains two large mixing angles and a third angle that is not exceedingly small. We cannot, therefore, associate any particular state  $|\nu_i\rangle$  with any particular lepton label  $e, \mu$  or  $\tau$ . Nevertheless, note that in the standard labeling the  $|\nu_1\rangle$  has the largest  $|\nu_e\rangle$  component ( $\sim 2/3$ ),  $|\nu_2\rangle$  contains  $\sim 1/3$  of the  $|\nu_e\rangle$  component and  $|\nu_3\rangle$  contains only a small  $\sim 2.5\%$   $|\nu_e\rangle$  component.

Neutrinos are produced in weak decays with a definite lepton flavor, and are typically detected by the charged current weak interaction again associated with a specific lepton flavor. Hence, the listings for the neutrino mass that follow are separated into the three associated charged lepton categories. Other properties (mean lifetime, magnetic moment, charge and charge radius) are no longer separated this way. If needed, the associated lepton flavor is reported in the footnotes.

Measured quantities (mass-squared, magnetic moments, mean lifetimes, *etc.*) all depend upon the mixing parameters  $|U_{\ell i}|^2$ , but to some extent also on experimental conditions (*e.g.*, on energy resolution). Most of these observables, in particular mass-squared, cannot distinguish between Dirac and Majorana neutrinos, and are unaffected by  $CP$  phases.

Direct neutrino mass measurements are usually based on the analysis of the kinematics of charged particles (leptons, pions) emitted together with neutrinos (flavor states) in various weak decays. The most sensitive neutrino mass measurement to date, involving electron type antineutrinos, is based on fitting the shape of the beta spectrum. The quantity  $\langle m_\beta^2 \rangle = \sum_i |U_{ei}|^2 m_{\nu_i}^2$  is determined or constrained, where the sum is over all mass eigenvalues  $m_{\nu_i}$  that are too close together to be resolved experimentally. If the energy resolution is better than  $\Delta m_{ij}^2 \equiv m_{\nu_i}^2 - m_{\nu_j}^2$ , the corresponding heavier  $m_{\nu_i}$  and mixing

# Lepton Particle Listings

## Neutrino Properties

parameter could be determined by fitting the resulting spectral anomaly (step or kink).

A limit on  $\langle m_\beta^2 \rangle$  implies an upper limit on the minimum value  $m_{\min}^2$  of  $m_{\nu_i}^2$ , independent of the mixing parameters  $U_{ei}$ :  $m_{\min}^2 \leq \langle m_\beta^2 \rangle$ . However, if and when the value of  $\langle m_\beta^2 \rangle$  is determined then its combination with the results derived from neutrino oscillations that give us the values of the neutrino mass-squared differences  $\Delta m_{ij}^2 \equiv m_i^2 - m_j^2$  and the mixing parameters  $|U_{ei}|^2$ , the individual neutrino mass squares  $m_{\nu_j}^2 = \langle m_\beta^2 \rangle - \sum_i |U_{ei}|^2 \Delta m_{ij}^2$  can be determined.

So far solar, reactor, atmospheric and accelerator neutrino oscillation experiments can be consistently described using three active neutrino flavors, i.e. two mass splittings and three mixing angles. However, several experiments with radioactive sources, reactors, and accelerators imply the possible existence of one or more non-interacting neutrino species that might be observable since they couple weakly to the flavor neutrinos  $|\nu_l\rangle$ .

Combined three neutrino analyses determine the squared mass differences and all three mixing angles to within reasonable accuracy. For given  $|\Delta m_{ij}^2|$  a limit on  $\langle m_\beta^2 \rangle$  from beta decay defines an upper limit on the maximum value  $m_{\max}$  of  $m_{\nu_i}$ :  $m_{\max}^2 \leq \langle m_\beta^2 \rangle + \sum_{i < j} |\Delta m_{ij}^2|$ . The analysis of the low energy beta decay of tritium, combined with the oscillation results, thus limits all active neutrino masses. Traditionally, experimental neutrino mass limits obtained from pion decay  $\pi^+ \rightarrow \mu^+ + \nu_\mu$  or the shape of the spectrum of decay products of the  $\tau$  lepton did not distinguish between flavor and mass eigenstates. These results are reported as limits of the  $\mu$  and  $\tau$  based neutrino mass. After the determination of the  $|\Delta m_{ij}^2|$ 's and the mixing angles  $\theta_{ij}$ , the corresponding neutrino mass limits are no longer competitive with those derived from low energy beta decays.

The spread of arrival times of the neutrinos from SN1987A, coupled with the measured neutrino energies, provided a time-of-flight limit on a quantity similar to  $\langle m_\beta \rangle \equiv \sqrt{\langle m_\beta^2 \rangle}$ . This statement, clothed in various degrees of sophistication, has been the basis for a very large number of papers. The resulting limits, however, are no longer comparable with the limits from tritium beta decay.

Constraint on the sum of the neutrino masses can be obtained from the analysis of the cosmic microwave background anisotropy, combined with the galaxy redshift surveys and other data. These limits are reported in a separate table (Sum of Neutrino Masses,  $m_{\text{tot}}$ ). Discussion concerning the model dependence of this limit is continuing.

### $\overline{\nu}$ MASS (electron based)

Those limits given below are for the square root of  $m_{\nu_e}^{2(\text{eff})} \equiv \sum_i |U_{ei}|^2 m_{\nu_i}^2$ . Limits that come from the kinematics of  ${}^3\text{H}\beta^- \overline{\nu}$  decay are the square roots of the limits for  $m_{\nu_e}^{2(\text{eff})}$ . Obtained from the measurements reported in the Listings for " $\overline{\nu}$  Mass Squared," below.

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2 OUR EVALUATION</b>				
<b>&lt; 2.05</b>	95	<sup>1</sup> ASEEV	11	SPEC ${}^3\text{H}\beta$ decay
<b>&lt; 2.3</b>	95	<sup>2</sup> KRAUS	05	SPEC ${}^3\text{H}\beta$ decay

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 5.8	95	<sup>3</sup> PAGLIAROLI	10	ASTR	SN1987A
< 21.7	90	<sup>4</sup> ARNABOLDI	03A	BOLO	<sup>187</sup> Re $\beta$ -decay
< 5.7	95	<sup>5</sup> LOREDO	02	ASTR	SN1987A
< 2.5	95	<sup>6</sup> LOBASHEV	99	SPEC	${}^3\text{H}\beta$ decay
< 2.8	95	<sup>7</sup> WEINHEIMER	99	SPEC	${}^3\text{H}\beta$ decay
< 4.35	95	<sup>8</sup> BELESEV	95	SPEC	${}^3\text{H}\beta$ decay
< 12.4	95	<sup>9</sup> CHING	95	SPEC	${}^3\text{H}\beta$ decay
< 92	95	<sup>10</sup> HIDDEMANN	95	SPEC	${}^3\text{H}\beta$ decay
15 $\pm$ $^{+32}_{-15}$		HIDDEMANN	95	SPEC	${}^3\text{H}\beta$ decay
< 19.6	95	KERNAN	95	ASTR	SN 1987A
< 7.0	95	<sup>11</sup> STOEFL	95	SPEC	${}^3\text{H}\beta$ decay
< 7.2	95	<sup>12</sup> WEINHEIMER	93	SPEC	${}^3\text{H}\beta$ decay
< 11.7	95	<sup>13</sup> HOLZSCHUH	92B	SPEC	${}^3\text{H}\beta$ decay
< 13.1	95	<sup>14</sup> KAWAKAMI	91	SPEC	${}^3\text{H}\beta$ decay
< 9.3	95	<sup>15</sup> ROBERTSON	91	SPEC	${}^3\text{H}\beta$ decay
< 14	95	AVIGNONE	90	ASTR	SN 1987A
< 16		SPERGEL	88	ASTR	SN 1987A
17 to 40		<sup>16</sup> BORIS	87	SPEC	${}^3\text{H}\beta$ decay

<sup>1</sup> ASEEV 11 report the analysis of the entire beta endpoint data, taken with the Troitsk integrating electrostatic spectrometer between 1997 and 2002 (some of the earlier runs were rejected), using a windowless gaseous tritium source. The fitted value of  $m_{\nu_e}$ , based on the method of Feldman and Cousins, is obtained from the upper limit of the fit for  $m_{\nu_e}^2$ . Previous analysis problems were resolved by careful monitoring of the tritium gas column density. Supersedes LOBASHEV 99 and BELESEV 95.

<sup>2</sup> KRAUS 05 is a continuation of the work reported in WEINHEIMER 99. This result represents the final analysis of data taken from 1997 to 2001. Various sources of systematic uncertainties have been identified and quantified. The background has been reduced compared to the initial running period. A spectral anomaly at the endpoint, reported in LOBASHEV 99, was not observed.

<sup>3</sup> PAGLIAROLI 10 is critical of the likelihood method used by LOREDO 02.

<sup>4</sup> ARNABOLDI 03A *et al.* report kinematical neutrino mass limit using  $\beta$ -decay of <sup>187</sup>Re. Bolometric AgReO<sub>4</sub> micro-calorimeters are used. Mass bound is substantially weaker than those derived from tritium  $\beta$ -decays but has different systematic uncertainties.

<sup>5</sup> LOREDO 02 updates LOREDO 89.

<sup>6</sup> LOBASHEV 99 report a new measurement which continues the work reported in BELESEV 95. This limit depends on phenomenological fit parameters used to derive their best fit to  $m_{\nu_e}^2$ , making unambiguous interpretation difficult. See the footnote under " $\overline{\nu}$  Mass Squared."

<sup>7</sup> WEINHEIMER 99 presents two analyses which exclude the spectral anomaly and result in an acceptable  $m_{\nu_e}^2$ . We report the most conservative limit, but the other is nearly the same. See the footnote under " $\overline{\nu}$  Mass Squared."

<sup>8</sup> BELESEV 95 (Moscow) use an integral electrostatic spectrometer with adiabatic magnetic collimation and a gaseous tritium sources. A fit to a normal Kurie plot above 18300–18350 eV (to avoid a low-energy anomaly) plus a monochromatic line 7–15 eV below the endpoint yields  $m_{\nu_e}^2 = -4.1 \pm 10.9 \text{ eV}^2$ , leading to this Bayesian limit.

<sup>9</sup> CHING 95 quotes results previously given by SUN 93; no experimental details are given. A possible explanation for consistently negative values of  $m_{\nu_e}^2$  is given.

<sup>10</sup> HIDDEMANN 95 (Munich) experiment uses atomic tritium embedded in a metal-dioxide lattice. Bayesian limit calculated from the weighted mean  $m_{\nu_e}^2 = 221 \pm 4244 \text{ eV}^2$  from the two runs listed below.

<sup>11</sup> STOEFL 95 (LLNL) result is the Bayesian limit obtained from the  $m_{\nu_e}^2$  errors given below but with  $m_{\nu_e}^2$  set equal to 0. The anomalous endpoint accumulation leads to a value of  $m_{\nu_e}^2$  which is negative by more than 5 standard deviations.

<sup>12</sup> WEINHEIMER 93 (Mainz) is a measurement of the endpoint of the tritium  $\beta$  spectrum using an electrostatic spectrometer with a magnetic guiding field. The source is molecular tritium frozen onto an aluminum substrate.

<sup>13</sup> HOLZSCHUH 92B (Zurich) result is obtained from the measurement  $m_{\nu_e}^2 = -24 \pm 48 \pm 61$  (1 $\sigma$  errors), in eV<sup>2</sup>, using the PDG prescription for conversion to a limit in  $m_{\nu_e}$ .

<sup>14</sup> KAWAKAMI 91 (Tokyo) experiment uses tritium-labeled arachidic acid. This result is the Bayesian limit obtained from the  $m_{\nu_e}^2$  limit with the errors combined in quadrature. This was also done in ROBERTSON 91, although the authors report a different procedure.

<sup>15</sup> ROBERTSON 91 (LANL) experiment uses gaseous molecular tritium. The result is in strong disagreement with the earlier claims by the ITP group [LUBIMOV 80, BORIS 87 (+ BORIS 88 erratum)] that  $m_{\nu_e}$  lies between 17 and 40 eV. However, the probability of a positive  $m_{\nu_e}^2$  is only 3% if statistical and systematic error are combined in quadrature.

<sup>16</sup> See also comment in BORIS 87B and erratum in BORIS 88.

### $\overline{\nu}$ MASS SQUARED (electron based)

Given troubling systematics which result in improbably negative estimators of  $m_{\nu_e}^{2(\text{eff})} \equiv \sum_i |U_{ei}|^2 m_{\nu_i}^2$ , in many experiments, we use only KRAUS 05 and LOBASHEV 99 for our average.

VALUE (eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<b>– 0.6 <math>\pm</math> 1.9 OUR AVERAGE</b>				
– 0.67 $\pm$ 2.53		<sup>1</sup> ASEEV	11	SPEC ${}^3\text{H}\beta$ decay
– 0.6 $\pm$ 2.2 $\pm$ 2.1		<sup>2</sup> KRAUS	05	SPEC ${}^3\text{H}\beta$ decay

See key on page 601

# Lepton Particle Listings

## Neutrino Properties

• • • We do not use the following data for averages, fits, limits, etc. • • •

– 1.9 ± 3.4 ± 2.2	3	LOBASHEV	99	SPEC	$^3\text{H}$ $\beta$ decay
– 3.7 ± 5.3 ± 2.1	4	WEINHEIMER	99	SPEC	$^3\text{H}$ $\beta$ decay
– 22 ± 4.8	5	BELESEV	95	SPEC	$^3\text{H}$ $\beta$ decay
129 ± 6010	6	HIDDEMANN	95	SPEC	$^3\text{H}$ $\beta$ decay
313 ± 5994	6	HIDDEMANN	95	SPEC	$^3\text{H}$ $\beta$ decay
– 130 ± 20 ± 15	7	STOEFL	95	SPEC	$^3\text{H}$ $\beta$ decay
– 31 ± 75 ± 48	8	SUN	93	SPEC	$^3\text{H}$ $\beta$ decay
– 39 ± 34 ± 15	9	WEINHEIMER	93	SPEC	$^3\text{H}$ $\beta$ decay
– 24 ± 48 ± 61	10	HOLZSCHUH	92B	SPEC	$^3\text{H}$ $\beta$ decay
– 65 ± 85 ± 65	11	KAWAKAMI	91	SPEC	$^3\text{H}$ $\beta$ decay
– 147 ± 68 ± 41	12	ROBERTSON	91	SPEC	$^3\text{H}$ $\beta$ decay

<sup>1</sup> ASEEV 11 report the analysis of the entire beta endpoint data, taken with the Troitsk integrating electrostatic spectrometer between 1997 and 2002, using a windowless gaseous tritium source. The analysis does not use the two additional fit parameters (see LOBASHEV 99) for a step-like structure near the endpoint. Using only the runs where the tritium gas column density was carefully monitored the need for such parameters was eliminated. Supersedes LOBASHEV 99 and BELESEV 95.

<sup>2</sup> KRAUS 05 is a continuation of the work reported in WEINHEIMER 99. This result represents the final analysis of data taken from 1997 to 2001. Problems with significantly negative squared neutrino masses, observed in some earlier experiments, have been resolved in this work.

<sup>3</sup> LOBASHEV 99 report a new measurement which continues the work reported in BELESEV 95. The data were corrected for electron trapping effects in the source, eliminating the dependence of the fitted neutrino mass on the fit interval. The analysis assuming a pure beta spectrum yields significantly negative fitted  $m_\nu^2 \approx -(20-10) \text{ eV}^2$ . This problem is attributed to a discrete spectral anomaly of about  $6 \times 10^{-11}$  intensity with a time-dependent energy of 5–15 eV below the endpoint. The data analysis accounts for this anomaly by introducing two extra phenomenological fit parameters resulting in a best fit of  $m_\nu^2 = -1.9 \pm 3.4 \pm 2.2 \text{ eV}^2$  which is used to derive a neutrino mass limit. However, the introduction of phenomenological fit parameters which are correlated with the derived  $m_\nu^2$  limit makes unambiguous interpretation of this result difficult.

<sup>4</sup> WEINHEIMER 99 is a continuation of the work reported in WEINHEIMER 93. Using a lower temperature of the frozen tritium source eliminated the dewetting of the  $T_2$  film, which introduced a dependence of the fitted neutrino mass on the fit interval in the earlier work. An indication for a spectral anomaly reported in LOBASHEV 99 has been seen, but its time dependence does not agree with LOBASHEV 99. Two analyses, which exclude the spectral anomaly either by choice of the analysis interval or by using a particular data set which does not exhibit the anomaly, result in acceptable  $m_\nu^2$  fits and are used to derive the neutrino mass limit published by the authors. We list the most conservative of the two.

<sup>5</sup> BELESEV 95 (Moscow) use an integral electrostatic spectrometer with adiabatic magnetic collimation and a gaseous tritium source. This value comes from a fit to a normal Kurie plot above 18300–18350 eV (to avoid a low-energy anomaly), including the effects of an apparent peak 7–15 eV below the endpoint.

<sup>6</sup> HIDDEMANN 95 (Munich) experiment uses atomic tritium embedded in a metal-dioxide lattice. They quote measurements from two data sets.

<sup>7</sup> STOEFL 95 (LLNL) uses a gaseous source of molecular tritium. An anomalous pileup of events at the endpoint leads to the negative value for  $m_\nu^2$ . The authors acknowledge that “the negative value for the best fit of  $m_\nu^2$  has no physical meaning” and discuss possible explanations for this effect.

<sup>8</sup> SUN 93 uses a tritiated hydrocarbon source. See also CHING 95.

<sup>9</sup> WEINHEIMER 93 (Mainz) is a measurement of the endpoint of the tritium  $\beta$  spectrum using an electrostatic spectrometer with a magnetic guiding field. The source is molecular tritium frozen onto an aluminum substrate.

<sup>10</sup> HOLZSCHUH 92B (Zurich) source is a monolayer of tritiated hydrocarbon.

<sup>11</sup> KAWAKAMI 91 (Tokyo) experiment uses tritium-labeled arachidic acid.

<sup>12</sup> ROBERTSON 91 (LANL) experiment uses gaseous molecular tritium. The result is in strong disagreement with the earlier claims by the ITEP group [LUBIMOV 80, BORIS 87 (+ BORIS 88 erratum)] that  $m_\nu$  lies between 17 and 40 eV. However, the probability of a positive  $m_\nu^2$  is only 3% if statistical and systematic error are combined in quadrature.

### $\nu$ MASS (electron based)

These are measurement of  $m_{\nu\mu}$  (in contrast to  $m_{\tau\mu}$  given above). The masses can be different for a Dirac neutrino in the absence of  $CPT$  invariance. The possible distinction between  $\nu$  and  $\bar{\nu}$  properties is usually ignored elsewhere in these Listings.

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<460	68	YASUMI	94	CNTR $^{163}\text{Ho}$ decay
<225	95	SPRINGER	87	CNTR $^{163}\text{Ho}$ decay

### $\nu$ MASS (muon based)

Limits given below are for the square root of  $m_{\nu\mu}^{2(\text{eff})} \equiv \sum_i |\mathcal{U}_{\mu i}|^2 m_{\nu_i}^2$ .

In some of the COSM papers listed below, the authors did not distinguish between weak and mass eigenstates.

OUR EVALUATION is based on OUR AVERAGE for the  $\pi^\pm$  mass and the ASSAMAGAN 96 value for the muon momentum for the  $\pi^\pm$  decay at rest. The limit is calculated using the unified classical analysis of FELDMAN 98 for a Gaussian distribution near a physical boundary. WARNING: since

$m_{\nu\mu}^{2(\text{eff})}$  is calculated from the differences of large numbers, it and the corresponding limits are extraordinarily sensitive to small changes in the pion mass, the decay muon momentum, and their errors. For example, the limits obtained using JECKELMANN 94, LENZ 98, and the weighted averages are 0.15, 0.29, and 0.19 MeV, respectively.

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.19 (CL = 90%) OUR EVALUATION</b>				
<0.17	90	<sup>1</sup> ASSAMAGAN	96	SPEC $m_\nu^2 = -0.016 \pm 0.023$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.15		<sup>2</sup> DOLGOV	95	COSM Nucleosynthesis
<0.48		<sup>3</sup> ENQVIST	93	COSM Nucleosynthesis
<0.3		<sup>4</sup> FULLER	91	COSM Nucleosynthesis
<0.42		<sup>4</sup> LAM	91	COSM Nucleosynthesis
<0.50	90	<sup>5</sup> ANDERHUB	82	SPEC $m_\nu^2 = -0.14 \pm 0.20$
<0.65	90	CLARK	74	ASP K $K_{\mu 3}$ decay

<sup>1</sup> ASSAMAGAN 96 measurement of  $p_\mu$  from  $\pi^+ \rightarrow \mu^+ \nu$  at rest combined with JECKELMANN 94 Solution B pion mass yields  $m_\nu^2 = -0.016 \pm 0.023$  with corresponding Bayesian limit listed above. If Solution A is used,  $m_\nu^2 = -0.143 \pm 0.024 \text{ MeV}^2$ . Replaces ASSAMAGAN 94.

<sup>2</sup> DOLGOV 95 removes earlier assumptions (DOLGOV 93) about thermal equilibrium below  $T_{\text{QCD}}$  for wrong-helicity Dirac neutrinos (ENQVIST 93, FULLER 91) to set more stringent limits.

<sup>3</sup> ENQVIST 93 bases limit on the fact that thermalized wrong-helicity Dirac neutrinos would speed up expansion of early universe, thus reducing the primordial abundance. FULLER 91 exploits the same mechanism but in the older calculation obtains a larger production rate for these states, and hence a lower limit. Neutrino lifetime assumed to exceed nucleosynthesis time,  $\sim 1 \text{ s}$ .

<sup>4</sup> Assumes neutrino lifetime  $> 1 \text{ s}$ . For Dirac neutrinos only. See also ENQVIST 93.

<sup>5</sup> ANDERHUB 82 kinematics is insensitive to the pion mass.

### $\nu$ MASS (tau based)

The limits given below are the square roots of limits for  $m_{\nu\tau}^{2(\text{eff})} \equiv \sum_i |\mathcal{U}_{\tau i}|^2 m_{\nu_i}^2$ .

In some of the ASTR and COSM papers listed below, the authors did not distinguish between weak and mass eigenstates.

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< 18.2	95		<sup>1</sup> BARATE	98F	ALEP 1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 28	95		<sup>2</sup> ATHANAS	00	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
< 27.6	95		<sup>3</sup> ACKERSTAFF	98T	OPAL 1990–1995 LEP runs
< 30	95	473	<sup>4</sup> AMMAR	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
< 60	95		<sup>5</sup> ANASTASSOV	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
< 0.37 or > 22			<sup>6</sup> FIELDS	97	COSM Nucleosynthesis
< 68	95		<sup>7</sup> SWAIN	97	THEO $m_\tau, \tau_\tau, \tau$ partial widths
< 29.9	95		<sup>8</sup> ALEXANDER	96M	OPAL 1990–1994 LEP runs
< 149			<sup>9</sup> BOTTINO	96	THEO $\pi, \mu, \tau$ leptonic decays
< 1 or > 25			<sup>10</sup> HANNESTAD	96C	COSM Nucleosynthesis
< 71	95		<sup>11</sup> SOBIE	96	THEO $m_\tau, \tau_\tau, B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$
< 24	95	25	<sup>12</sup> BUSKULIC	95H	ALEP 1991–1993 LEP runs
< 0.19			<sup>13</sup> DOLGOV	95	COSM Nucleosynthesis
< 3			<sup>14</sup> SIGL	95	ASTR SN 1987A
< 0.4 or > 30			<sup>15</sup> DODELSON	94	COSM Nucleosynthesis
< 0.1 or > 50			<sup>16</sup> KAWASAKI	94	COSM Nucleosynthesis
155–225			<sup>17</sup> PERES	94	THEO $\pi, K, \mu, \tau$ weak decays
< 32.6	95	113	<sup>18</sup> CINABRO	93	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
< 0.3 or > 35			<sup>19</sup> DOLGOV	93	COSM Nucleosynthesis
< 0.74			<sup>20</sup> ENQVIST	93	COSM Nucleosynthesis
< 31	95	19	<sup>21</sup> ALBRECHT	92M	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
< 0.3			<sup>22</sup> FULLER	91	COSM Nucleosynthesis
< 0.5 or > 25			<sup>23</sup> KOLB	91	COSM Nucleosynthesis
< 0.42			<sup>22</sup> LAM	91	COSM Nucleosynthesis

<sup>1</sup> BARATE 98F result based on kinematics of  $2939 \tau^- \rightarrow 2\pi^- \pi^+ \nu_\tau$  and  $52 \tau^- \rightarrow 3\pi^- 2\pi^+ (\pi^0) \nu_\tau$  decays. If possible 2.5% excited  $a_1$  decay is included in 3-prong sample analysis, limit increases to 19.2 MeV.

<sup>2</sup> ATHANAS 00 bound comes from analysis of  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \pi^0 \nu_\tau$  decays.

<sup>3</sup> ACKERSTAFF 98T use  $\tau^- \rightarrow 5\pi^\pm \nu_\tau$  decays to obtain a limit of 43.2 MeV (95%CL). They combine this with ALEXANDER 96M value using  $\tau^- \rightarrow 3h^\pm \nu_\tau$  decays to obtain quoted limit.

<sup>4</sup> AMMAR 98 limit comes from analysis of  $\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau$  and  $\tau^- \rightarrow 2\pi^- \pi^+ 2\pi^0 \nu_\tau$  decay modes.

<sup>5</sup> ANASTASSOV 97 derive limit by comparing their  $m_\tau$  measurement (which depends on  $m_{\nu_\tau}$ ) to BAI 96  $m_\tau$  threshold measurement.

# Lepton Particle Listings

## Neutrino Properties

- <sup>6</sup> FIELDS 97 limit for a Dirac neutrino. For a Majorana neutrino the mass region  $< 0.93$  or  $> 31$  MeV is excluded. These bounds assume  $N_\nu < 4$  from nucleosynthesis; a wider excluded region occurs with a smaller  $N_\nu$  upper limit.
- <sup>7</sup> SWAIN 97 derive their limit from the Standard Model relationships between the tau mass, lifetime, branching fractions for  $\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau$ ,  $\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau$ ,  $\tau^- \rightarrow \pi^- \nu_\tau$ , and  $\tau^- \rightarrow K^- \nu_\tau$ , and the muon mass and lifetime by assuming lepton universality and using world average values. Limit is reduced to 48 MeV when the CLEO  $\tau$  mass measurement (BALEST 93) is included; see CLEO's more recent  $m_{\nu_\tau}$  limit (ANASTASSOV 97). Consideration of mixing with a fourth generation heavy neutrino yields  $\sin^2 \theta_L < 0.016$  (95%CL).
- <sup>8</sup> ALEXANDER 96M bound comes from analyses of  $\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau$  and  $\tau^- \rightarrow h^- h^- h^+ \nu_\tau$  decays.
- <sup>9</sup> BOTTINO 96 assumes three generations of neutrinos with mixing, finds consistency with massless neutrinos with no mixing based on 1995 data for masses, lifetimes, and leptonic partial widths.
- <sup>10</sup> HANNESTAD 96C limit is on the mass of a Majorana neutrino. This bound assumes  $N_\nu < 4$  from nucleosynthesis. A wider excluded region occurs with a smaller  $N_\nu$  upper limit. This paper is the corrected version of HANNESTAD 96; see the erratum: HANNESTAD 96B.
- <sup>11</sup> SOBIE 96 derive their limit from the Standard Model relationship between the tau mass, lifetime, and leptonic branching fraction, and the muon mass and lifetime, by assuming lepton universality and using world average values.
- <sup>12</sup> BUSKULIC 95H bound comes from a two-dimensional fit of the visible energy and invariant mass distribution of  $\tau \rightarrow 5\pi(\pi^0)\nu_\tau$  decays. Replaced by BARATE 98F.
- <sup>13</sup> DOLGOV 95 removes earlier assumptions (DOLGOV 93) about thermal equilibrium below  $T_{\text{QCD}}$  for wrong-helicity Dirac neutrinos (ENQVIST 93, FULLER 91) to set more stringent limits. DOLGOV 96 argues that a possible window near 20 MeV is excluded.
- <sup>14</sup> SIGL 95 exclude massive Dirac or Majorana neutrinos with lifetimes between  $10^{-3}$  and  $10^8$  seconds if the decay products are predominantly  $\gamma$  or  $e^+e^-$ .
- <sup>15</sup> DODELSON 94 calculate constraints on  $\nu_\tau$  mass and lifetime from nucleosynthesis for 4 generic decay modes. Limits depend strongly on decay mode. Quoted limit is valid for all decay modes of Majorana neutrinos with lifetime greater than about 300s. For Dirac neutrinos limits change to  $< 0.3$  or  $> 33$ .
- <sup>16</sup> KAWASAKI 94 excluded region is for Majorana neutrino with lifetime  $> 1000$  s. Other limits are given as a function of  $\nu_\tau$  lifetime for decays of the type  $\nu_\tau \rightarrow \nu_\mu \phi$  where  $\phi$  is a Nambu-Goldstone boson.
- <sup>17</sup> PERES 94 used PDG 92 values for parameters to obtain a value consistent with mixing. Reexamination by BOTTINO 96 which included radiative corrections and 1995 PDG parameters resulted in two allowed regions,  $m_3 < 70$  MeV and  $140 \text{ MeV} < m_3 < 149$  MeV.
- <sup>18</sup> CINABRO 93 bound comes from analysis of  $\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau$  and  $\tau^- \rightarrow 2\pi^- \pi^+ 2\pi^0 \nu_\tau$  decay modes.
- <sup>19</sup> DOLGOV 93 assumes neutrino lifetime  $> 100$  s. For Majorana neutrinos, the low mass limit is 0.5 MeV. KAWANO 92 points out that these bounds can be overcome for a Dirac neutrino if it possesses a magnetic moment. See also DOLGOV 96.
- <sup>20</sup> ENQVIST 93 bases limit on the fact that thermalized wrong-helicity Dirac neutrinos would speed up expansion of early universe, thus reducing the primordial abundance. FULLER 91 exploits the same mechanism but in the older calculation obtains a larger production rate for these states, and hence a lower limit. Neutrino lifetime assumed to exceed nucleosynthesis time,  $\sim 1$  s.
- <sup>21</sup> ALBRECHT 92M reports measurement of a slightly lower  $\tau$  mass, which has the effect of reducing the  $\nu_\tau$  mass reported in ALBRECHT 88B. Bound is from analysis of  $\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau$  mode.
- <sup>22</sup> Assumes neutrino lifetime  $> 1$  s. For Dirac neutrinos. See also ENQVIST 93.
- <sup>23</sup> KOLB 91 exclusion region is for Dirac neutrino with lifetime  $> 1$  s; other limits are given.

## SUM OF NEUTRINO MASSES

Revised January 2016 by K.A. Olive (University of Minnesota).

The limits on low mass ( $m_\nu \lesssim 1$  MeV) neutrinos apply to  $m_{\text{tot}}$  given by

$$m_{\text{tot}} = \sum_\nu (g_\nu/2) m_\nu,$$

where  $g_\nu$  is the number of spin degrees of freedom for  $\nu$  plus  $\bar{\nu}$ :  $g_\nu = 4$  for neutrinos with Dirac masses;  $g_\nu = 2$  for Majorana neutrinos. Stable neutrinos in this mass range make a contribution to the total energy density of the Universe which is given by

$$\rho_\nu = m_{\text{tot}} n_\nu = m_{\text{tot}} (3/11) n_\gamma,$$

where the factor 3/11 is the ratio of (light) neutrinos to photons. Writing  $\Omega_\nu = \rho_\nu/\rho_c$ , where  $\rho_c$  is the critical energy density of the Universe, and using  $n_\gamma = 412 \text{ cm}^{-3}$ , we have

$$\Omega_\nu h^2 = m_{\text{tot}} / (94 \text{ eV}).$$

While an upper limit to the matter density of  $\Omega_m h^2 < 0.12$  would constrain  $m_{\text{tot}} < 11$  eV, much stronger constraints are obtained from a combination of observations of the CMB, the amplitude of density fluctuations on smaller scales from the clustering of galaxies and the Lyman- $\alpha$  forest, baryon acoustic oscillations, and new Hubble parameter data. These combine to give an upper limit of around 0.2 eV, and may, in the near future, be able to provide a lower bound on the sum of the neutrino masses.

## SUM OF THE NEUTRINO MASSES, $m_{\text{tot}}$

(Defined in the above note), of effectively stable neutrinos (i.e., those with mean lives greater than or equal to the age of the universe). These papers assumed Dirac neutrinos. When necessary, we have generalized the results reported so they apply to  $m_{\text{tot}}$ . For other limits, see SZALAY 76, VYSOTSKY 77, BERNSTEIN 81, FREESE 84, SCHRAMM 84, and COWSIK 85.

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.15$	95	<sup>1</sup> PALANQUE-...	15	COSM SDSS/BOSS
$< 0.12$	95	<sup>2</sup> PALANQUE-...	15A	COSM SDSS/BOSS
$< 0.23$	95	<sup>3</sup> ADE	14	COSM Planck
$0.320 \pm 0.081$		<sup>4</sup> BATTYE	14	COSM
$0.35 \pm 0.10$		<sup>5</sup> BEUTLER	14	COSM BOSS
$0.22 \pm 0.09$ $-0.10$		<sup>6</sup> COSTANZI	14	COSM
$< 0.22$	95	<sup>7</sup> GIUSARMA	14	COSM
$0.32 \pm 0.11$		<sup>8</sup> HOU	14	COSM
$< 0.26$	95	<sup>9</sup> LEISTEDT	14	COSM
$< 0.18$	95	<sup>10</sup> RIEMER-SOR...	14	COSM
$< 0.24$	68	<sup>11</sup> MORESCO	12	COSM
$< 0.29$	95	<sup>12</sup> XIA	12	COSM
$< 0.81$	95	<sup>13</sup> SAITO	11	COSM SDSS
$< 0.44$	95	<sup>14</sup> HANNESTAD	10	COSM
$< 0.6$	95	<sup>15</sup> SEKIGUCHI	10	COSM
$< 0.28$	95	<sup>16</sup> THOMAS	10	COSM
$< 1.1$		<sup>17</sup> ICHIKI	09	COSM
$< 1.3$	95	<sup>18</sup> KOMATSU	09	COSM WMAP
$< 1.2$		<sup>19</sup> TERENO	09	COSM
$< 0.33$		<sup>20</sup> VIKHLININ	09	COSM
$< 0.28$		<sup>21</sup> BERNARDIS	08	COSM
$< 0.17-2.3$		<sup>22</sup> FOGLI	07	COSM
$< 0.42$	95	<sup>23</sup> KRISTIANSEN	07	COSM
$< 0.63-2.2$		<sup>24</sup> ZUNCKEL	07	COSM
$< 0.24$	95	<sup>25</sup> CIRELLI	06	COSM
$< 0.62$	95	<sup>26</sup> HANNESTAD	06	COSM
$< 1.2$		<sup>27</sup> SANCHEZ	06	COSM
$< 0.17$	95	<sup>28</sup> SELJAK	06	COSM
$< 2.0$	95	<sup>29</sup> ICHIKAWA	05	COSM
$< 0.75$		<sup>30</sup> BARGER	04	COSM
$< 1.0$		<sup>31</sup> CROTTY	04	COSM
$< 0.7$		<sup>32</sup> SPERGEL	03	COSM WMAP
$< 0.9$		<sup>33</sup> LEWIS	02	COSM
$< 4.2$		<sup>34</sup> WANG	02	COSM CMB
$< 2.7$		<sup>35</sup> FUKUGITA	00	COSM
$< 5.5$		<sup>36</sup> CROFT	99	ASTR Ly $\alpha$ power spec
$< 180$		SZALAY	74	COSM
$< 132$		COWSIK	72	COSM
$< 280$		MARX	72	COSM
$< 400$		GRSHTEIN	66	COSM

<sup>1</sup> Constrains the total mass of neutrinos using the Lyman  $\alpha$  forest power spectrum obtained by BOSS. The analysis includes CMB data from Planck, WMAP, ACT, and SPT. Limit improves to 0.14 when BAO data are included. Superseded by PALANQUE-DELABROUILLE 15A.

<sup>2</sup> Constrains the total mass of neutrinos using the Lyman- $\alpha$  forest power spectrum obtained by BOSS. The analysis includes CMB data from Planck, ACT, and SPT. Limit is unchanged when BAO data are included. Supersedes PALANQUE-DELABROUILLE 15.

<sup>3</sup> Constrains the total mass of neutrinos from Planck CMB data along with WMAP polarization, high L, and BAO data.

<sup>4</sup> Finite neutrino mass fit to resolve discrepancy between CMB and lensing measurements.

<sup>5</sup> Fit to the total mass of neutrinos from BOSS data along with WMAP CMB data and data from other BAO constraints and weak lensing.

<sup>6</sup> Fit to the total mass of neutrinos from Planck CMB data along with BAO.

<sup>7</sup> Constrains the total mass of neutrinos from Planck CMB data combined with baryon acoustic oscillation data from BOSS and HST data on the Hubble parameter.

<sup>8</sup> Fit based on the SPT-SZ survey combined with CMB, BAO, and  $H_0$  data.

<sup>9</sup> Constrains the total mass of neutrinos (marginalizing over the effective number of neutrino species) from CMB, CMB lensing, BAO, and galaxy clustering data.



# Lepton Particle Listings

## Neutrino Properties

- <sup>10</sup> Constrains the total mass of neutrinos from Planck CMB data combined with baryon acoustic oscillation data from BOSS, 6dFGS, SDSS, WiggleZ data on the galaxy power spectrum, and HST data on the Hubble parameter. The limit is increased to 0.25 eV if a lower bound to the sum of neutrino masses of 0.04 eV is assumed.
- <sup>11</sup> Constrains the total mass of neutrinos from observational Hubble parameter data with seven-year WMAP data and the most recent estimate of  $H_0$ .
- <sup>12</sup> Constrains the total mass of neutrinos from the CFHTLS combined with seven-year WMAP data and a prior on the Hubble parameter. Limit is relaxed to 0.41 eV when small scales affected by non-linearities are removed.
- <sup>13</sup> Constrains the total mass of neutrinos from the Sloan Digital Sky Survey and the five-year WMAP data.
- <sup>14</sup> Constrains the total mass of neutrinos from the 7-year WMAP data including SDSS and HST data. Limit relaxes to 1.19 eV when CMB data is used alone. Supersedes HANNESTAD 06.
- <sup>15</sup> Constrains the total mass of neutrinos from a combination of CMB data, a recent measurement of  $H_0$  (SHOES), and baryon acoustic oscillation data from SDSS.
- <sup>16</sup> Constrains the total mass of neutrinos from SDSS MegaZ LRG DR7 galaxy clustering data combined with CMB, HST, supernovae and baryon acoustic oscillation data. Limit relaxes to 0.47 eV when the equation of state parameter,  $w \neq -1$ .
- <sup>17</sup> Constrains the total mass of neutrinos from weak lensing measurements when combined with CMB. Limit improves to 0.54 eV when supernovae and baryon acoustic oscillation observations are included. Assumes  $\Lambda$ CDM model.
- <sup>18</sup> Constrains the total mass of neutrinos from five-year WMAP data. Limit improves to 0.67 eV when supernovae and baryon acoustic oscillation observations are included. Limits quoted assume the  $\Lambda$ CDM model. Supersedes SPERGEL 07.
- <sup>19</sup> Constrains the total mass of neutrinos from weak lensing measurements when combined with CMB. Limit improves to  $0.03 < \Sigma m_\nu < 0.54$  eV when supernovae and baryon acoustic oscillation observations are included. The slight preference for massive neutrinos at the two-sigma level disappears when systematic errors are taken into account. Assumes  $\Lambda$ CDM model.
- <sup>20</sup> Constrains the total mass of neutrinos from recent Chandra X-ray observations of galaxy clusters when combined with CMB, supernovae, and baryon acoustic oscillation measurements. Assumes flat universe and constant dark-energy equation of state,  $w$ .
- <sup>21</sup> Constraints the total mass of neutrinos from recent CMB and SOSS LRG power spectrum data along with bias mass relations from SDSS, DEEP2, and Lyman-Break Galaxies. It assumes  $\Lambda$ CDM model. Limit degrades to 0.59 eV in a more general  $w$ CDM model.
- <sup>22</sup> Constrains the total mass of neutrinos from neutrino oscillation experiments and cosmological data. The most conservative limit uses only WMAP three-year data, while the most stringent limit includes CMB, large-scale structure, supernova, and Lyman-alpha data.
- <sup>23</sup> Constrains the total mass of neutrinos from recent CMB, large scale structure, SNIa, and baryon acoustic oscillation data. The limit relaxes to 1.75 when WMAP data alone is used with no prior. Paper shows results with several combinations of data sets. Supersedes KRISTIANSEN 06.
- <sup>24</sup> Constrains the total mass of neutrinos from the CMB and the large scale structure data. The most conservative limit is obtained when generic initial conditions are allowed.
- <sup>25</sup> Constrains the total mass of neutrinos from recent CMB, large scale structure, Lyman-alpha forest, and SNIa data.
- <sup>26</sup> Constrains the total mass of neutrinos from recent CMB and large scale structure data. See also GOOBAR 06. Superseded by HANNESTAD 10.
- <sup>27</sup> Constrains the total mass of neutrinos from the CMB and the final 2dF Galaxy Redshift Survey.
- <sup>28</sup> Constrains the total mass of neutrinos from the CMB experiments alone, assuming  $\Lambda$ CDM Universe. FUKUGITA 06 show that this result is unchanged by the 3-year WMAP data.
- <sup>29</sup> Constrains the total mass of neutrinos from the power spectrum of fluctuations derived from the Sloan Digital Sky Survey and the 2dF galaxy redshift survey, WMAP and 27 other CMB experiments and measurements by the HST Key project.
- <sup>30</sup> Constrains the total mass of neutrinos from the power spectrum of fluctuations derived from the Sloan Digital Sky Survey, the 2dF galaxy redshift survey, WMAP and ACBAR. The limit is strengthened to 0.6 eV when measurements by the HST Key project and supernovae data are included.
- <sup>31</sup> Constrains the fractional contribution of neutrinos to the total matter density in the Universe from WMAP data combined with other CMB measurements, the 2dFGRS data, and Lyman  $\alpha$  data. The limit does not noticeably change if the Lyman  $\alpha$  data are not used.
- <sup>32</sup> LEWIS 02 constrains the total mass of neutrinos from the power spectrum of fluctuations derived from the CMB, HST Key project, 2dF galaxy redshift survey, supernovae type Ia, and BBN.
- <sup>33</sup> WANG 02 constrains the total mass of neutrinos from the power spectrum of fluctuations derived from the CMB and other cosmological data sets such as galaxy clustering and the Lyman  $\alpha$  forest.
- <sup>34</sup> FUKUGITA 00 is a limit on neutrino masses from structure formation. The constraint is based on the clustering scale  $\sigma_8$  and the COBE normalization and leads to a conservative limit of 0.9 eV assuming 3 nearly degenerate neutrinos. The quoted limit is on the sum of the light neutrino masses.
- <sup>35</sup> CROFT 99 result based on the power spectrum of the Ly  $\alpha$  forest. If  $\Omega_{\text{matter}} < 0.5$ , the limit is improved to  $m_\nu < 2.4$  ( $\Omega_{\text{matter}}/0.17-1$ ) eV.

### Limits on MASSES of Light Stable Right-Handed $\nu$ (with necessarily suppressed interaction strengths)

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<100–200	<sup>1</sup> OLIVE	82	COSM Dirac $\nu$
<200–2000	<sup>1</sup> OLIVE	82	COSM Majorana $\nu$

<sup>1</sup> Depending on interaction strength  $G_R$  where  $G_R < G_F$ .

### Limits on MASSES of Heavy Stable Right-Handed $\nu$ (with necessarily suppressed interaction strengths)

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
> 10	<sup>1</sup> OLIVE	82	COSM $G_R/G_F < 0.1$
>100	<sup>1</sup> OLIVE	82	COSM $G_R/G_F < 0.01$

<sup>1</sup> These results apply to heavy Majorana neutrinos and are summarized by the equation:  $m_\nu > 1.2$  GeV ( $G_F/G_R$ ). The bound saturates, and if  $G_R$  is too small no mass range is allowed.

### $\nu$ CHARGE

VALUE (units: electron charge)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.1 $\times 10^{-12}$	90	<sup>1</sup> CHEN	14A	TEXO Nuclear reactor
<1.5 $\times 10^{-12}$	90	<sup>2</sup> STUDENIKIN	14	Nuclear reactor
<3.7 $\times 10^{-12}$	90	<sup>3</sup> GNINENKO	07	RVUE Nuclear reactor
<2 $\times 10^{-14}$		<sup>4</sup> RAFFELT	99	ASTR Red giant luminosity
<6 $\times 10^{-14}$		<sup>5</sup> RAFFELT	99	ASTR Solar cooling
<4 $\times 10^{-4}$		<sup>6</sup> BABU	94	RVUE BEBC beam dump
<3 $\times 10^{-4}$		<sup>7</sup> DAVIDSON	91	RVUE SLAC $e^-$ beam dump
<2 $\times 10^{-15}$		<sup>8</sup> BARBIELLINI	87	ASTR SN 1987A
<1 $\times 10^{-13}$		<sup>9</sup> BERNSTEIN	63	ASTR Solar energy losses

<sup>1</sup> CHEN 14A use the Multi-Configuration RPPA method to analyze reactor  $\bar{\nu}_e$  scattering on Ge atoms with 300 eV recoil energy threshold to obtain this limit.

<sup>2</sup> STUDENIKIN 14 uses the limit on  $\mu_\nu$  from BEDA 13 and the 2.8 keV threshold of the electron recoil energy to obtain this limit.

<sup>3</sup> GNINENKO 07 use limit on  $\bar{\nu}_e$  magnetic moment from Li 03b to derive this result. The limit is considerably weaker than the limits on the charge of  $\nu_e$  and  $\bar{\nu}_e$  from various astrophysics considerations.

<sup>4</sup> This RAFFELT 99 limit applies to all neutrino flavors which are light enough (<5 keV) to be emitted from globular-cluster red giants.

<sup>5</sup> This RAFFELT 99 limit is derived from the helioseismological limit on a new energy-loss channel of the Sun, and applies to all neutrino flavors which are light enough (<1 keV) to be emitted from the sun.

<sup>6</sup> BABU 94 use COOPER-SARKAR 92 limit on  $\nu$  magnetic moment to derive quoted result. It applies to  $\nu_\tau$ .

<sup>7</sup> DAVIDSON 91 use data from early SLAC electron beam dump experiment to derive charge limit as a function of neutrino mass. It applies to  $\nu_\tau$ .

<sup>8</sup> Exact BARBIELLINI 87 limit depends on assumptions about the intergalactic or galactic magnetic fields and about the direct distance and time through the field. It applies to  $\nu_e$ .

<sup>9</sup> The limit applies to all flavors.

### $\nu$ (MEAN LIFE) / MASS

Measures  $[\sum |U_{\ell j}|^2 \Gamma_j m_j]^{-1}$ , where the sum is over mass eigenstates which cannot be resolved experimentally. Some of the limits constrain the radiative decay and are based on the limit of the corresponding photon flux. Other apply to the decay of a heavier neutrino into the lighter one and a Majoron or other invisible particle. Many of these limits apply to any  $\nu$  within the indicated mass range.

Limits on the radiative decay are either directly based on the limits of the corresponding photon flux, or are derived from the limits on the neutrino magnetic moments. In the later case the transition rate for  $\nu_i \rightarrow \nu_j + \gamma$

is constrained by  $\Gamma_{ij} = \frac{1}{\tau_{ij}} = \frac{(m_i^2 - m_j^2)^3}{m_i^3} \mu_{ij}^2$  where  $\mu_{ij}$  is the neutrino transition moment in the mass eigenstates basis. Typically, the limits on lifetime based on the magnetic moments are many orders of magnitude more restrictive than limits based on the nonobservation of photons.

VALUE (s/eV)	CL%	DOCUMENT ID	TECN	COMMENT
> 15.4	90	<sup>1</sup> KRAKAUER	91	CNTR $\nu_\mu, \bar{\nu}_\mu$ at LAMPF
> 7 $\times 10^9$		<sup>2</sup> RAFFELT	85	ASTR
> 300	90	<sup>3</sup> REINES	74	CNTR $\bar{\nu}_e$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 10 <sup>5</sup> –10 <sup>10</sup>	95	<sup>4</sup> CECCHINI	11	ASTR $\nu_2 \rightarrow \nu_1$ radiative decay
	90	<sup>5</sup> MIRIZZI	07	CMB radiative decay
	90	<sup>6</sup> MIRIZZI	07	CIB radiative decay
		<sup>7</sup> WONG	07	CNTR Reactor $\bar{\nu}_e$
> 0.11	90	<sup>8</sup> XIN	05	CNTR Reactor $\nu_e$
		<sup>9</sup> XIN	05	CNTR Reactor $\nu_e$
> 0.004	90	<sup>10</sup> AHARMIM	04	SNO quasidegen. $\nu$ masses
> 4.4 $\times 10^{-5}$	90	<sup>10</sup> AHARMIM	04	SNO hierarchical $\nu$ masses
$\gtrsim 100$	95	<sup>11</sup> CECCHINI	04	ASTR Radiative decay for $\nu$ mass > 0.01 eV
> 0.067	90	<sup>12</sup> EGUCHI	04	KLND quasidegen. $\nu$ masses
> 1.1 $\times 10^{-3}$	90	<sup>12</sup> EGUCHI	04	KLND hierarchical $\nu$ masses
> 8.7 $\times 10^{-5}$	99	<sup>13</sup> BANDYOPA...	03	FIT nonradiative decay
$\geq 4200$	90	<sup>14</sup> DERBIN	02B	CNTR Solar $pp$ and Be $\nu$
> 2.8 $\times 10^{-5}$	99	<sup>15</sup> JOSHUPURA	02B	FIT nonradiative decay

# Lepton Particle Listings

## Neutrino Properties

		16 DOLGOV	99 COSM
		17 BILLER	98 ASTR $m_{\nu}=0.05\text{--}1\text{ eV}$
		18,19 BLUDMAN	92 ASTR $m_{\nu}<50\text{ eV}$
		20 DODELSON	92 ASTR $m_{\nu}=1\text{--}300\text{ keV}$
		20 DODELSON	92 ASTR $m_{\nu}=1\text{--}300\text{ keV}$
		21 GRANEK	91 COSM Decaying $l^0$
		22 KRAKAUER	91 CNTR $\nu_e$ at LAMPF
		23 WALKER	90 ASTR $m_{\nu}=0.03\text{--}\sim 2\text{ MeV}$
		19,24 CHUPP	89 ASTR $m_{\nu}<20\text{ eV}$
		19 KOLB	89 ASTR $m_{\nu}<20\text{ eV}$
		25 RAFFELT	89 RVUE $\overline{\nu}$ (Dirac, Majorana)
		26 RAFFELT	89B ASTR
		27 VONFEILIT...	88 ASTR
		28 OBERAUER	87 $\overline{\nu}_R$ (Dirac)
		28 OBERAUER	87 $\overline{\nu}$ (Majorana)
		28 OBERAUER	87 $\overline{\nu}_L$ (Dirac)
		28 KETOV	86 CNTR $\overline{\nu}$ (Dirac)
		28 KETOV	86 CNTR $\overline{\nu}$ (Majorana)
		29 BINETRUY	84 COSM $m_{\nu}\sim 1\text{ MeV}$
		30 FRANK	81 CNTR $\nu\overline{\nu}$ LAMPF
		31 STECKER	80 ASTR $m_{\nu}=10\text{--}100\text{ eV}$
		30 BLIETSCHAU	78 HLBC $\nu_{\mu}$ , CERN GGM
		30 BLIETSCHAU	78 HLBC $\overline{\nu}_{\mu}$ , CERN GGM
		32 FALK	77 ASTR $m_{\nu}<10\text{ MeV}$
		30 BARNES	77 DBC $\nu$ , ANL 12-ft
		33 COWSIK	77 ASTR
		30 BELLOTTI	78 HLBC $\nu$ , CERN GGM
		30 BELLOTTI	76 HLBC $\overline{\nu}$ , CERN GGM

- 1 KRAKAUER 91 quotes the limit  $\tau/m_{\nu_1} > (0.75a^2 + 21.65a + 26.3)\text{ s/eV}$ , where  $a$  is a parameter describing the asymmetry in the neutrino decay defined as  $dN_{\nu}/dcos\theta = (1/2)(1 + a\cos\theta)$ . The parameter  $a=0$  for a Majorana neutrino, but can vary from  $-1$  to  $1$  for a Dirac neutrino. The bound given by the authors is the most conservative (which applies for  $a = -1$ ).
- 2 RAFFELT 85 limit on the radiative decay is from solar  $x$ - and  $\gamma$ -ray fluxes. Limit depends on  $\nu$  flux from  $pp$ , now established from GALLEX and SAGE to be  $> 0.5$  of expectation.
- 3 REINES 74 looked for  $\nu$  of nonzero mass decaying radiatively to a neutral of lesser mass  $+\gamma$ . Used liquid scintillator detector near fission reactor. Finds lab lifetime  $6 \times 10^7\text{ s}$  or more. Above value of (mean life)/mass assumes average effective neutrino energy of  $0.2\text{ MeV}$ . To obtain the limit  $6 \times 10^7\text{ s}$  REINES 74 assumed that the full  $\overline{\nu}_e$  reactor flux could be responsible for yielding decays with photon energies in the interval  $0.1\text{ MeV} - 0.5\text{ MeV}$ . This represents some overestimate so their lower limit is an over-estimate of the lab lifetime (VOGEL 84). If so, OBERAUER 87 may be comparable or better.
- 4 CECCHINI 11 search for radiative decays of solar neutrinos into visible photons during the 2006 total solar eclipse. The range of (mean life)/mass values corresponds to a range of  $\nu_1$  masses between  $10^{-4}$  and  $0.1\text{ eV}$ .
- 5 MIRIZZI 07 determine a limit on the neutrino radiative decay from analysis of the maximum allowed distortion of the CMB spectrum as measured by the COBE/FIRAS. For the decay  $\nu_2 \rightarrow \nu_1$  the lifetime limit is  $\lesssim 4 \times 10^{20}\text{ s}$  for  $m_{\min} \lesssim 0.14\text{ eV}$ . For transition with the  $|\Delta m_{31}|$  mass difference the lifetime limit is  $\sim 2 \times 10^{19}\text{ s}$  for  $m_{\min} \lesssim 0.14\text{ eV}$  and  $\sim 5 \times 10^{20}\text{ s}$  for  $m_{\min} \gtrsim 0.14\text{ eV}$ .
- 6 MIRIZZI 07 determine a limit on the neutrino radiative decay from analysis of the cosmic infrared background (CIB) using the Spitzer Observatory data. For transition with the  $|\Delta m_{31}|$  mass difference they obtain the lifetime limit  $\sim 10^{20}\text{ s}$  for  $m_{\min} \lesssim 0.14\text{ eV}$ .
- 7 WONG 07 use their limit on the neutrino magnetic moment together with the assumed experimental value of  $\Delta m_{13}^2 \sim 2 \times 10^{-3}\text{ eV}^2$  to obtain  $\tau_{13}/m_1^3 > 3.2 \times 10^{27}\text{ s/eV}^3$  for the radiative decay in the case of the inverted mass hierarchy. Similarly to RAFFELT 89 this limit can be violated if electric and magnetic moments are equal to each other. Analogous, but numerically somewhat different limits are obtained for  $\tau_{23}$  and  $\tau_{21}$ .
- 8 XIN 05 search for the  $\gamma$  from radiative decay of  $\nu_e$  produced by the electron capture on  $^{51}\text{Cr}$ . No events were seen and the limit on  $\tau/m_{\nu}$  was derived. This is a weaker limit on the decay of  $\nu_e$  than KRAKAUER 91.
- 9 XIN 05 use their limit on the neutrino magnetic moment of  $\nu_e$  together with the assumed experimental value of  $\Delta m_{13}^2 \sim 2 \times 10^{-3}\text{ eV}^2$  to obtain  $\tau_{13}/m_1^3 > 1 \times 10^{23}\text{ s/eV}^3$  for the radiative decay in the case of the inverted mass hierarchy. Similarly to RAFFELT 89 this limit can be violated if electric and magnetic moments are equal to each other. Analogous, but numerically somewhat different limits are obtained for  $\tau_{23}$  and  $\tau_{21}$ . Again, this limit is specific for  $\nu_e$ .
- 10 AHARMIM 04 obtained these results from the solar  $\overline{\nu}_e$  flux limit set by the SNO measurement assuming  $\nu_2$  decay through nonradiative process  $\nu_2 \rightarrow \overline{\nu}_1 X$ , where  $X$  is a Majoron or other invisible particle. Limits are given for the cases of quasidegenerate and hierarchical neutrino masses.
- 11 CECCHINI 04 obtained this bound through the observations performed on the occasion of the 21 June 2001 total solar eclipse, looking for visible photons from radiative decays of solar neutrinos. Limit is a  $\tau/m_{\nu_2}$  in  $\nu_2 \rightarrow \nu_1 \gamma$ . Limit ranges from  $\sim 100$  to  $10^7\text{ s/eV}$  for  $0.01 < m_{\nu_1} < 0.1\text{ eV}$ .
- 12 EGUCHI 04 obtained these results from the solar  $\overline{\nu}_e$  flux limit set by the KamLAND measurement assuming  $\nu_2$  decay through nonradiative process  $\nu_2 \rightarrow \overline{\nu}_1 X$ , where  $X$  is a Majoron or other invisible particle. Limits are given for the cases of quasidegenerate and hierarchical neutrino masses.
- 13 The ratio of the lifetime over the mass derived by BANDYOPADHYAY 03 is for  $\nu_2$ . They obtained this result using the following solar-neutrino data: total rates measured in Cl and Ga experiments, the Super-Kamiokande's zenith-angle spectra, and SNO's day and night spectra. They assumed that  $\nu_1$  is the lowest mass, stable or nearly stable neutrino

state and  $\nu_2$  decays through nonradiative Majoron emission process,  $\nu_2 \rightarrow \overline{\nu}_1 + J$ , or through nonradiative process with all the final state particles being sterile. The best fit is obtained in the region of the LMA solution.

- 14 DERBIN 02b (also BACK 03b) obtained this bound for the radiative decay from the results of background measurements with Counting Test Facility (the prototype of the Borexino detector). The laboratory gamma spectrum is given as  $dN_{\gamma}/dcos\theta = (1/2)(1 + \alpha\cos\theta)$  with  $\alpha=0$  for a Majorana neutrino, and  $\alpha$  varying to  $-1$  to  $1$  for a Dirac neutrino. The listed bound is for the case of  $\alpha=0$ . The most conservative bound  $1.5 \times 10^{13}\text{ s eV}^{-1}$  is obtained for the case of  $\alpha=-1$ .
- 15 The ratio of the lifetime over the mass derived by JOSHIPURA 02b is for  $\nu_2$ . They obtained this result from the total rates measured in all solar neutrino experiments. They assumed that  $\nu_1$  is the lowest mass, stable or nearly stable neutrino state and  $\nu_2$  decays through nonradiative process like Majoron emission decay,  $\nu_2 \rightarrow \nu_1' + J$  where  $\nu_1'$  state is sterile. The exact limit depends on the specific solution of the solar neutrino problem. The quoted limit is for the LMA solution.
- 16 DOLGOV 99 places limits in the (Majorana)  $\tau$ -associated  $\nu$  mass-lifetime plane based on nucleosynthesis. Results would be considerably modified if neutrino oscillations exist.
- 17 BILLER 98 use the observed TeV  $\gamma$ -ray spectra to set limits on the mean life of any radiatively decaying neutrino between  $0.05$  and  $1\text{ eV}$ . Curve shows  $\tau_{\nu}/B_{\gamma} > 0.15 \times 10^{21}\text{ s}$  at  $0.05\text{ eV}$ ,  $> 1.2 \times 10^{21}\text{ s}$  at  $0.17\text{ eV}$ ,  $> 3 \times 10^{21}\text{ s}$  at  $1\text{ eV}$ , where  $B_{\gamma}$  is the branching ratio to photons.
- 18 BLUDMAN 92 sets additional limits by this method for higher mass ranges. Cosmological limits are also obtained.
- 19 Limit on the radiative decay based on nonobservation of  $\gamma$ 's in coincidence with  $\nu$ 's from SN 1987A.
- 20 DODELSON 92 range is for wrong-helicity keV mass Dirac  $\nu$ 's from the core of neutron star in SN 1987A decaying to  $\nu$ 's that would have interacted in KAM2 or LMB detectors.
- 21 GRANEK 91 considers heavy neutrino decays to  $\gamma\nu_L$  and  $3\nu_L$ , where  $m_{\nu_L} < 100\text{ keV}$ . Lifetime is calculated as a function of heavy neutrino mass, branching ratio into  $\gamma\nu_L$ , and  $m_{\nu_L}$ .
- 22 KRAKAUER 91 quotes the limit for  $\nu_e$ ,  $\tau/m_{\nu} > (0.3a^2 + 9.8a + 15.9)\text{ s/eV}$ , where  $a$  is a parameter describing the asymmetry in the radiative neutrino decay defined as  $dN_{\nu}/dcos\theta = (1/2)(1 + a\cos\theta)$   $a=0$  for a Majorana neutrino, but can vary from  $-1$  to  $1$  for a Dirac neutrino. The bound given by the authors is the most conservative (which applies for  $a = -1$ ).
- 23 WALKER 90 uses SN 1987A  $\gamma$  flux limits after 289 days.
- 24 CHUPP 89 should be multiplied by a branching ratio (about 1) and a detection efficiency (about 1/4), and pertains to radiative decay of any neutrino to a lighter or sterile neutrino.
- 25 RAFFELT 89 uses KYULDJIEV 84 to obtain  $\tau m^3 > 3 \times 10^{18}\text{ s eV}^3$  (based on  $\overline{\nu}_e e^-$  cross sections). The bound for the radiative decay is not valid if electric and magnetic transition moments are equal for Dirac neutrinos.
- 26 RAFFELT 89b analyze stellar evolution and exclude the region  $3 \times 10^{12} < \tau m^3 < 3 \times 10^{21}\text{ s eV}^3$ .
- 27 Model-dependent theoretical analysis of SN 1987A neutrinos. Quoted limit is for  $[\sum_j |U_{\ell j}|^2 \Gamma_j m_j]^{-1}$ , where  $\ell=\mu, \tau$ . Limit is  $3.3 \times 10^{14}\text{ s/eV}$  for  $\ell=e$ .
- 28 OBERAUER 87 looks for photons and  $e^+e^-$  pairs from radiative decays of reactor neutrinos.
- 29 BINETRUY 84 finds  $\tau < 10^8\text{ s}$  for neutrinos in a radiation-dominated universe.
- 30 These experiments look for  $\nu_k \rightarrow \nu_j \gamma$  or  $\overline{\nu}_k \rightarrow \overline{\nu}_j \gamma$ .
- 31 STECKER 80 limit based on UV background; result given is  $\tau > 4 \times 10^{22}\text{ s}$  at  $m_{\nu}=20\text{ eV}$ .
- 32 FALK 78 finds lifetime constraints based on supernova energetics.
- 33 COWSIK 77 considers variety of scenarios. For neutrinos produced in the big bang, present limits on optical photon flux require  $\tau > 10^{23}\text{ s}$  for  $m_{\nu} \sim 1\text{ eV}$ . See also COWSIK 79 and GOLDMAN 79.

## $\nu$ MAGNETIC MOMENT

The coupling of neutrinos to an electromagnetic field is a characterized by a  $3 \times 3$  matrix  $\lambda$  of the magnetic ( $\mu$ ) and electric ( $d$ ) dipole moments ( $\lambda = \mu - id$ ). For Majorana neutrinos the matrix  $\lambda$  is antisymmetric and only transition moments are allowed, while for Dirac neutrinos  $\lambda$  is a general  $3 \times 3$  matrix. In the standard electroweak theory extended to include neutrino masses (see FUJIKAWA 80)  $\mu_{\nu} = 3eG_F m_{\nu}/(8\pi^2 \sqrt{2}) = 3.2 \times 10^{-19}(m_{\nu}/\text{eV})\mu_B$ , i.e. it is unobservably small given the known small neutrino masses. In more general models there is no longer a proportionality between neutrino mass and its magnetic moment, even though only massive neutrinos have nonvanishing magnetic moments without fine tuning.

Laboratory bounds on  $\lambda$  are obtained via elastic  $\nu$ - $e$  scattering, where the scattered neutrino is not observed. The combinations of matrix elements of  $\lambda$  that are constrained by various experiments depend on the initial neutrino flavor and on its propagation between source and detector (e.g., solar  $\nu_e$  and reactor  $\overline{\nu}_e$  do not constrain the same combinations). The listings below therefore identify the initial neutrino flavor.

Other limits, e.g. from various stellar cooling processes, apply to all neutrino flavors. Analogous flavor independent, but weaker, limits are obtained from the analysis of  $e^+e^- \rightarrow \nu\overline{\nu}\gamma$  collider experiments.

See key on page 601

# Lepton Particle Listings

## Neutrino Properties

VALUE ( $10^{-10} \mu_B$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>0.29</b>	90	1 BEDA 13	CNTR	Reactor $\bar{\nu}_e$
< <b>6.8</b>	90	2 AUERBACH 01	LSND	$\nu_e e$ , $\nu_\mu e$ scattering
< <b>3900</b>	90	3 SCHWIENHO...01	DONU	$\nu_\tau e^- \rightarrow \nu_\tau e^-$
< 0.022	90	4 ARCEO-DIAZ 15	ASTR	Red giants
< 0.1	95	5 CORSICO 14	ASTR	
< 0.05	95	6 MILLER-BER...14B	ASTR	
< 0.045	95	7 VIAUX 13A	ASTR	Globular cluster M5
< 0.32	90	8 BEDA 10	CNTR	Reactor $\bar{\nu}_e$
< 2.2	90	9 DENIZ 10	TEXO	Reactor $\bar{\nu}_e$
< 0.011–0.027	90	10 KUZNETSOV 09	ASTR	$\nu_L \rightarrow \nu_R$ in SN1987A
< 0.54	90	11 ARPESELLA 08A	BORX	Solar $\nu$ spectrum shape
< 0.58	90	12 BEDA 07	CNTR	Reactor $\bar{\nu}_e$
< 0.74	90	13 WONG 07	CNTR	Reactor $\bar{\nu}_e$
< 0.9	90	14 DARAKTCHIEVA...05		Reactor $\bar{\nu}_e$
< 130	90	15 XIN 05	CNTR	Reactor $\nu_e$
< 37	95	16 GRIFOLS 04	FIT	Solar $^8\text{B}$ $\nu$ (SNO NC)
< 3.6	90	17 LIU 04	SKAM	Solar $\nu$ spectrum shape
< 1.1	90	18 LIU 04	SKAM	Solar $\nu$ spectrum shape (LMA region)
< 5.5	90	19 BACK 03B	CNTR	Solar $pp$ and Be $\nu$
< 1.0	90	20 DARAKTCHIEVA...03		Reactor $\bar{\nu}_e$
< 1.3	90	21 LI 03B	CNTR	Reactor $\bar{\nu}_e$
< 2	90	22 GRIMUS 02	FIT	solar + reactor (Majorana $\nu$ )
< 80000	90	23 TANIMOTO 00	RVUE	$e^+ e^- \rightarrow \nu \bar{\nu}$
< 0.01–0.04	90	24 AYALA 99	ASTR	$\nu_L \rightarrow \nu_R$ in SN1987A
< 1.5	90	25 BEACOM 99	SKAM	$\nu$ spectrum shape
< 0.03	90	26 RAFFELT 99	ASTR	Red giant luminosity
< 4	90	27 RAFFELT 99	ASTR	Solar cooling
< 44000	90	28 ABREU 97J	DLPH	$e^+ e^- \rightarrow \nu \bar{\nu} \gamma$ at LEP
< 33000	90	29 ACCIARRI 97Q	L3	$e^+ e^- \rightarrow \nu \bar{\nu} \gamma$ at LEP
< 0.62	90	29 ELMFORS 97	COSM	Depolarization in early universe plasma
< 27000	95	30 ESCRIBANO 97	RVUE	$\Gamma(Z \rightarrow \nu \bar{\nu})$ at LEP
< 30	90	31 VILAIN 95B	CHM2	$\nu_\mu e \rightarrow \nu_\mu e$
< 55000	90	32 GOULD 94	RVUE	$e^+ e^- \rightarrow \nu \bar{\nu} \gamma$ at LEP
< 1.9	95	33 DERBIN 93	CNTR	Reactor $\bar{\nu}_e \rightarrow \bar{\nu}_e$
< 5400	90	34 COOPER...92	BEBC	$\nu_\tau e^- \rightarrow \nu_\tau e^-$
< 2.4	90	35 VIDYAKIN 92	CNTR	Reactor $\bar{\nu}_e \rightarrow \bar{\nu}_e$
< 56000	90	36 DESHPANDE 91	RVUE	$e^+ e^- \rightarrow \nu \bar{\nu} \gamma$
< 100	95	37 DORENBOSCH...91	CHRM	$\nu_\mu e \rightarrow \nu_\mu e$
< 8.5	90	38 AHRENS 90	CNTR	$\nu_\mu e \rightarrow \nu_\mu e$
< 10.8	90	39 KRAKAUER 90	CNTR	LAMPF $\nu e \rightarrow \nu e$
< 7.4	90	40 KRAKAUER 90	CNTR	LAMPF $(\nu_\mu, \bar{\nu}_\mu) e$ elast.
< 0.02	90	41 RAFFELT 90	ASTR	Red giant luminosity
< 0.1	90	42 RAFFELT 89B	ASTR	Cooling helium stars
< 40000	90	43 FUKUGITA 88	COSM	Primordial magn. fields
< .3	90	44 GROTH 88	RVUE	$e^+ e^- \rightarrow \nu \bar{\nu} \gamma$
< 0.11	90	45 RAFFELT 88B	ASTR	He burning stars
< 0.0006	90	46 FUKUGITA 87	ASTR	Cooling helium stars
< 0.1–0.2	90	47 NUSSINOV 87	ASTR	Cosmic EM back-grounds
< 0.85	90	48 MORGAN 81	COSM	$^4\text{He}$ abundance
< 0.6	90	49 BEG 78	ASTR	Stellar plasmons
< 81	90	50 SUTHERLAND 76	ASTR	Red giants + degenerate dwarfs
< 1	90	51 KIM 74	RVUE	$\bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$
< 14	90	52 BERNSTEIN 63	ASTR	Solar cooling
< 14	90	53 COWAN 57	CNTR	Reactor $\bar{\nu}$

- 1 BEDA 13 report  $\bar{\nu}_e e^-$  scattering results, using the Kalinin Nuclear Power Plant and a shielded Ge detector. The recoil electron spectrum is analyzed between 2.5 and 55 keV. Supersedes BEDA 07. Supersedes BEDA 10. This is the most stringent limit on the magnetic moment of reactor  $\bar{\nu}_e$ .
- 2 AUERBACH 01 limit is based on the LSND  $\nu_e$  and  $\nu_\mu$  electron scattering measurements. The limit is slightly more stringent than KRAKAUER 90.
- 3 SCHWIENHORST 01 quote an experimental sensitivity of  $4.9 \times 10^{-7}$ .
- 4 ARCEO-DIAZ 15 constrains the neutrino magnetic moment from observation of the tip of the red giant branch in the globular cluster  $\omega$ -Centauri.
- 5 CORSICO 14 constrains the neutrino magnetic moment from observations of white dwarf pulsations.
- 6 MILLER-BERTOLAMI 14B constrains the neutrino magnetic moment from observations of the white dwarf luminosity function of the Galactic disk.
- 7 VIAUX 13A constrains the neutrino magnetic moment from observations of the globular cluster M5.
- 8 BEDA 10 report  $\bar{\nu}_e e^-$  scattering results, using the Kalinin Nuclear Power Plant and a shielded Ge detector. The recoil electron spectrum is analyzed between 2.9 and 45 keV. Supersedes BEDA 07. Supersedes by BEDA 13.
- 9 DENIZ 10 observe reactor  $\bar{\nu}_e e^-$  scattering with recoil kinetic energies 3–8 MeV using CsI(Tl) detectors. The observed rate and spectral shape are consistent with the Standard Model prediction, leading to the reported constraint on  $\bar{\nu}_e$  magnetic moment.

- 10 KUZNETSOV 09 obtain a limit on the flavor averaged magnetic moment of Dirac neutrinos from the time averaged neutrino signal of SN1987A. Improves and supersedes the analysis of BARBIERI 88 and AYALA 99.
- 11 ARPESELLA 08A obtained this limit using the shape of the recoil electron energy spectrum from the Borexino 192 live days of solar neutrino data.
- 12 BEDA 07 performed search for electromagnetic  $\bar{\nu}_e e^-$  scattering at Kalininskaya nuclear reactor. A Ge detector with active and passive shield was used and the electron recoil spectrum between 3.0 and 61.3 keV analyzed. Superseded by BEDA 10.
- 13 WONG 07 performed search for non-standard  $\bar{\nu}_e e^-$  scattering at the Kuo-Sheng nuclear reactor. Ge detector equipped with active anti-Compton shield is used. Most stringent laboratory limit on magnetic moment of reactor  $\bar{\nu}_e$ . Supersedes LI 03B.
- 14 DARAKTCHIEVA 05 present the final analysis of the search for non-standard  $\bar{\nu}_e e^-$  scattering component at Bugey nuclear reactor. Full kinematical event reconstruction of both the kinetic energy above 700 keV and scattering angle of the recoil electron, by use of TPC. Most stringent laboratory limit on magnetic moment. Supersedes DARAKTCHIEVA 03.
- 15 XIN 05 evaluated the  $\nu_e$  flux at the Kuo-Sheng nuclear reactor and searched for non-standard  $\nu_e e^-$  scattering. Ge detector equipped with active anti-Compton shield was used. This laboratory limit on magnetic moment is considerably less stringent than the limits for reactor  $\bar{\nu}_e$ , but is specific to  $\nu_e$ .
- 16 GRIFOLS 04 obtained this bound using the SNO data of the solar  $^8\text{B}$  neutrino flux measured with deuteron breakup. This bound applies to  $\mu_{\text{eff}} = (\mu_{21}^2 + \mu_{22}^2 + \mu_{23}^2)^{1/2}$ .
- 17 LIU 04 obtained this limit using the shape of the recoil electron energy spectrum from the Super-Kamiokande-I 1496 days of solar neutrino data. Neutrinos are assumed to have only diagonal magnetic moments,  $\mu_{\nu 1} = \mu_{\nu 2}$ . This limit corresponds to the oscillation parameters in the vacuum oscillation region.
- 18 LIU 04 obtained this limit using the shape of the recoil electron energy spectrum from the Super-Kamiokande-I 1496 live-day solar neutrino data, by limiting the oscillation parameter region in the LMA region allowed by solar neutrino experiments plus KamLAND.  $\mu_{\nu 1} = \mu_{\nu 2}$  is assumed. In the LMA region, the same limit would be obtained even if neutrinos have off-diagonal magnetic moments.
- 19 BACK 03B obtained this bound from the results of background measurements with Counting Test Facility (the prototype of the Borexino detector). Standard Solar Model flux was assumed. This  $\mu_\nu$  can be different from the reactor  $\mu_\nu$  in certain oscillation scenarios (see BEACOM 99).
- 20 DARAKTCHIEVA 03 searched for non-standard  $\bar{\nu}_e e^-$  scattering component at Bugey nuclear reactor. Full kinematical event reconstruction by use of TPC. Superseded by DARAKTCHIEVA 05.
- 21 LI 03B used Ge detector in active shield near nuclear reactor to test for nonstandard  $\bar{\nu}_e e^-$  scattering.
- 22 GRIMUS 02 obtain stringent bounds on all Majorana neutrino transition moments from a simultaneous fit of LMA-MSW oscillation parameters and transition moments to global solar neutrino data + reactor data. Using only solar neutrino data, a 90% CL bound of  $6.3 \times 10^{-10} \mu_B$  is obtained.
- 23 TANIMOTO 00 combined  $e^+ e^- \rightarrow \nu \bar{\nu} \gamma$  data from VENUS, TOPAZ, and AMY.
- 24 AYALA 99 improves the limit of BARBIERI 88.
- 25 BEACOM 99 obtain the limit using the shape, but not the absolute magnitude which is affected by oscillations, of the solar neutrino spectrum obtained by Superkamiokande (825 days). This  $\mu_\nu$  can be different from the reactor  $\mu_\nu$  in certain oscillation scenarios.
- 26 RAFFELT 99 is an update of RAFFELT 90. This limit applies to all neutrino flavors which are light enough ( $< 5$  keV) to be emitted from globular-cluster red giants. This limit pertains equally to electric dipole moments and magnetic transition moments, and it applies to both Dirac and Majorana neutrinos.
- 27 RAFFELT 99 is essentially an update of BERNSTEIN 63, but is derived from the helioseismological limit on a new energy-loss channel of the Sun. This limit applies to all neutrino flavors which are light enough ( $< 1$  keV) to be emitted from the Sun. This limit pertains equally to electric dipole and magnetic transition moments, and it applies to both Dirac and Majorana neutrinos.
- 28 ACCIARRI 97Q result applies to both direct and transition magnetic moments and for  $q^2=0$ .
- 29 ELMFORS 97 calculate the rate of depolarization in a plasma for neutrinos with a magnetic moment and use the constraints from a big-bang nucleosynthesis on additional degrees of freedom.
- 30 Applies to absolute value of magnetic moment.
- 31 DERBIN 93 determine the cross section for 0.6–2.0 MeV electron energy as  $(1.28 \pm 0.63) \times \sigma_{\text{weak}}$ . However, the (reactor on – reactor off)/(reactor off) is only  $\sim 1/100$ .
- 32 COOPER-SARKAR 92 assume  $f_{D_s}/f_\pi = 2$  and  $D_s, \bar{D}_s$  production cross section =  $2.6 \mu\text{b}$  to calculate  $\nu$  flux.
- 33 VIDYAKIN 92 limit is from a  $e\bar{\nu}_e$  elastic scattering experiment. No experimental details are given except for the cross section from which this limit is derived. Signal/noise was 1/10. The limit uses  $\sin^2 \theta_W = 0.23$  as input.
- 34 DORENBOSCH 91 corrects an incorrect statement in DORENBOSCH 89 that the  $\nu$  magnetic moment is  $< 1 \times 10^{-9}$  at the 95%CL. DORENBOSCH 89 measures both  $\nu_\mu e$  and  $\bar{\nu}_e e$  elastic scattering and assume  $\mu(\nu) = \mu(\bar{\nu})$ .
- 35 KRAKAUER 90 experiment fully reported in ALLEN 93.
- 36 RAFFELT 90 limit applies for a diagonal magnetic moment of a Dirac neutrino, or for a transition magnetic moment of a Majorana neutrino. In the latter case, the same analysis gives  $< 1.4 \times 10^{-12}$ . Limit at 95%CL obtained from  $\delta M_C$ .
- 37 Significant dependence on details of stellar models.
- 38 FUKUGITA 88 find magnetic dipole moments of any two neutrino species are bounded by  $\mu < 10^{-16} [10^{-9} G/B_0]$  where  $B_0$  is the present-day intergalactic field strength.
- 39 GROTH 88 combined data from MAC, ASP, CELLO, and Mark J.

# Lepton Particle Listings

## Neutrino Properties

- <sup>40</sup> For  $m_\nu = 8\text{--}200$  eV. NUSSINOV 87 examines transition magnetic moments for  $\nu_\mu \rightarrow \nu_e$  and obtain  $< 3 \times 10^{-15}$  for  $m_\nu > 16$  eV and  $< 6 \times 10^{-14}$  for  $m_\nu > 4$  eV.
- <sup>41</sup> We obtain above limit from SUTHERLAND 76 using their limit  $f < 1/3$ .
- <sup>42</sup> KIM 74 is a theoretical analysis of  $\bar{\nu}_\mu$  reaction data.

### NEUTRINO CHARGE RADIUS SQUARED

We report limits on the so-called neutrino charge radius squared. While the straight-forward definition of a neutrino charge radius has been proven to be gauge-dependent and, hence, unphysical (LEE 77c), there have been recent attempts to define a physically observable neutrino charge radius (BERNABEU 00, BERNABEU 02). The issue is still controversial (FUJIKAWA 03, BERNABEU 03). A more general interpretation of the experimental results is that they are limits on certain nonstandard contributions to neutrino scattering.

VALUE ( $10^{-32}$ cm <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<b>-2.1 to 3.3</b>	90	<sup>1</sup> DENIZ	10 TEXO	Reactor $\bar{\nu}_e e$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.53 to 0.68	90	<sup>2</sup> HIRSCH	03	$\nu_\mu e$ scat.
-8.2 to 9.9	90	<sup>3</sup> HIRSCH	03	anomalous $e^+ e^- \rightarrow \nu \bar{\nu} \gamma$
-2.97 to 4.14	90	<sup>4</sup> AUERBACH	01 LSND	$\nu_e e \rightarrow \nu_e e$
-0.6 to 0.6	90	VILAIN	95B CHM2	$\nu_\mu e$ elastic scat.
0 ± 2.7		ALLEN	93 CNTR	LAMPPF $\nu e \rightarrow \nu e$
< 2.3	95	MOURAO	92 ASTR	HOME/KAM2 $\nu$ rates
< 7.3	90	<sup>5</sup> VIDYAKIN	92 CNTR	Reactor $\bar{\nu} e \rightarrow \bar{\nu} e$
1.1 ± 2.3		ALLEN	91 CNTR	Repl. by ALLEN 93
-1.1 ± 1.0		<sup>6</sup> AHRENS	90 CNTR	$\nu_\mu e$ elastic scat.
-0.3 ± 1.5		<sup>6</sup> DORENBOS...	89 CHRM	$\nu_\mu e$ elastic scat.
		<sup>7</sup> GRIFOLS	89B ASTR	SN 1987A

- <sup>1</sup> DENIZ 10 observe reactor  $\bar{\nu}_e e$  scattering with recoil kinetic energies 3–8 MeV using CsI(Tl) detectors. The observed rate and spectral shape are consistent with the Standard Model prediction, leading to the reported constraint on  $\bar{\nu}_e$  charge radius.
- <sup>2</sup> Based on analysis of CCFR 98 results. Limit is on  $\langle r_V^2 \rangle + \langle r_A^2 \rangle$ . The CHARM II and E734 at BNL results are reanalyzed, and weaker bounds on the charge radius squared than previously published are obtained. The NuTeV result is discussed; when tentatively interpreted as  $\nu_\mu$  charge radius it implies  $\langle r_V^2 \rangle + \langle r_A^2 \rangle = (4.20 \pm 1.64) \times 10^{-33}$  cm<sup>2</sup>.
- <sup>3</sup> Results of LEP-2 are interpreted as limits on the axial-vector charge radius squared of a Majorana  $\nu_\mu$ . Slightly weaker limits for both vector and axial-vector charge radius squared are obtained for the Dirac case, and somewhat weaker limits are obtained from the analysis of lower energy data (LEP-1.5 and TRISTAN).
- <sup>4</sup> AUERBACH 01 measure  $\nu_e e$  elastic scattering with LSND detector. The cross section agrees with the Standard Model expectation, including the charge and neutral current interference. The 90% CL applies to the range shown.
- <sup>5</sup> VIDYAKIN 92 limit is from a  $e \bar{\nu}$  elastic scattering experiment. No experimental details are given except for the cross section from which this limit is derived. Signal/noise was 1/10. The limit uses  $\sin^2 \theta_W = 0.23$  as input.
- <sup>6</sup> Result is obtained from reanalysis given in ALLEN 91, followed by our reduction to obtain  $1\sigma$  errors.
- <sup>7</sup> GRIFOLS 89b sets a limit of  $\langle r^2 \rangle < 0.2 \times 10^{-32}$  cm<sup>2</sup> for right-handed neutrinos.

### REFERENCES FOR Neutrino Properties

ARCEO-DIAZ 15 ASP 70 1 S. Arceo-Diaz *et al.*

PALANQUE... 15 JCAP 1502 045 N. Palanque-Desabrouille *et al.*

PALANQUE... 15A JCAP 1511 011 N. Palanque-Desabrouille *et al.*

ADE 14 AA 571 A16 P.A.R. Ade *et al.* (Planck Collab.)

BATTYE 14 PRL 112 051303 R.A. Battye, A. Moss (MCHS, NOTT)

BEUTLER 14 MNRAS 444 3501 F. Beutler *et al.* (BOSS Collab.)

CHEN 14A PR D90 011301 J.-W. Chen *et al.* (TEXONO Collab.)

CORSICO 14 JCAP 1408 054 A.H. Corsico

COSTANZI 14 JCAP 1410 081 M. Costanzi *et al.* (TRST, TRSTI)

GIUSARMA 14 PR D90 043507 E. Giusarma *et al.*

HOU 14 APJ 782 74 Z. Hou *et al.*

LEISTEDT 14 PRL 113 041301 B. Leistedt, H.V. Peiris, L. Verde

MILLER-BER... 14B AA 562 A123 M.M. Miller Bertolami (MPiG, LAPL)

RIEMER-SØK... 14 PR D89 103505 S. Riemer-Sørensen, D. Parkinson, T. M. Davis

STUDENIKIN 14 EPL 107 21001 A.I. Studenikin

BEDA 13 PPNL 10 139 A.G. Beda *et al.* (GEMMA Collab.)

VIAUX 13A PRL 111 231301 N. Viaux *et al.*

MORESCO 12 JCAP 1207 053 M. Moresco *et al.*

XIA 12 JCAP 1206 010 J.-Q. Xia *et al.*

ASEEV 11 PR D84 112003 V.N. Aseev *et al.*

CECCHINI 11 ASP 34 486 S. Cecchini *et al.*

SAITO 11 PR D83 043529 S. Saito, M. Takada, A. Taruya

BEDA 10 PPNL 7 406 A.G. Beda *et al.* (GEMMA Collab.)

DENIZ 10 PR D81 072001 M. Deniz *et al.* (TEXONO Collab.)

HANNESTAD 10 JCAP 1008 001 S. Hannestad *et al.*

PAGLIAROLI 10 ASP 33 287 G. Pagliaroli, F. Rossi-Torres, E. Vissani (INFN+)

SEKIGUCHI 10 JCAP 1003 015 T. Sekiguchi *et al.*

THOMAS 10 PRL 105 031301 S.A. Thomas, F.B. Abdalla, O. Lahav (LOUC)

ICHIKI 09 PR D79 023520 K. Ichiki, M. Takada, T. Takahashi

KOMATSU 09 APJS 180 330 E. Komatsu *et al.*

KUZNETSOV 09 IJMP A24 5977 A.V. Kuznetsov, N.V. Mikheev, A.A. Okrugin (YARO)

TERENO 09 AA 500 657 I. Tereno *et al.*

VIKHLININ 09 APJ 692 1060 A. Vikhlinin *et al.*

ARPESELLA 08A PRL 101 091302 C. Arpessella *et al.* (Borexino Collab.)

BERNARDIS 08 PR D78 083535 F. De Bernardis *et al.*

BEDA 07 PAN 70 1873 A.G. Beda *et al.*

Translated from YAF 70 1925.

FOGLI 07 PR D75 053001 G.L. Fogli *et al.*

GNINENKO 07 PR D75 075014 S.N. Gninenko, N.V. Krasnikov, A. Rubbia

KRISTIANSEN 07 PR D75 083510 J. Kristiansen, O. Elgaroy, H. Dahle

MIRIZZI 07 PR D76 053007 A. Mirizzi, D. Montanino, P.D. Serpico

SPEGEL 07 APJS 170 377 D.N. Spergel *et al.*

WONG 07 PR D75 012001 H.T. Wong *et al.* (TEXONO Collab.)

ZUNCKEL 07 JCAP 0708 004 C. Zunckel, P. Ferreira

CIRELLI 06 JCAP 0612 013 M. Cirelli *et al.*

FUKUGITA 06 PR D74 027302 M. Fukugita *et al.*

GOOBAR 06 JCAP 0606 019 A. Goobar *et al.*

HANNESTAD 06 JCAP 0611 016 S. Hannestad, G. Raffelt

KRISTIANSEN 06 PR D74 123005 J. Kristiansen, O. Elgaroy, H. Eriksen

SANCHEZ 06 MNRAS 366 189 A.G. Sanchez *et al.*

SELIJAK 06 JCAP 0610 014 U. Seljak, A. Slosar, P. McDonald

DARAKTCH... 05 PL B615 153 Z. Daraktchieva *et al.* (MUNU Collab.)

ICHIKAWA 05 PR D71 043001 K. Ichikawa, M. Fukugita, M. Kawasaki (ICRR)

KRAUS 05 EPJ C40 447 Ch. Kraus *et al.*

XIN 05 PR D72 012006 B. Xin *et al.* (TEXONO Collab.)

AHARMIM 04 PR D70 093014 B. Aharmim *et al.* (SNO Collab.)

BARGER 04 PL B595 55 V. Barger, D. Marfatia, A. Tregre

CECCHINI 04 ASP 21 183 S. Cecchini *et al.* (BGNA+)

CROTTY 04 PR D69 123007 P. Crotty, J. Lesgourgues, S. Pastor

EGUCHI 04 PRL 92 071301 K. Eguchi *et al.* (KamLAND Collab.)

GRIFOLS 04 PL B587 184 J.A. Grifols, E. Masso, S. Mohanty (BARC, AHMED)

LIU 04 PRL 93 021802 D.W. Liu *et al.* (Super-Kamiokande Collab.)

ARNABOLDI 03A PRL 91 161802 C. Arnaboldi *et al.*

BACK 03B PL B563 35 H.O. Back *et al.* (Borexino Collab.)

BANDYOPAD... 03 PL B555 33 A. Bandyopadhyay, S. Choubey, S. Goswami (SAHA+)

BERNABEU 03 hep-ph/0303202 J. Bernabeu, J. Papavassiliou, J. Vidal

DARAKTCH... 03 PL B564 190 Z. Daraktchieva *et al.* (MUNU Collab.)

FUJIKAWA 03 hep-ph/0303188 K. Fujikawa, R. Shrock

HIRSCH 03 PR D67 033005 M. Hirsch, E. Nardi, D. Restrepo

LI 03B PRL 90 131802 H.B. Li *et al.* (TEXONO Collab.)

SPEGEL 03 APJS 148 175 D.N. Spergel *et al.*

BERNABEU 02 PRL 89 101802 J. Bernabeu, J. Papavassiliou, J. Vidal

Also PRL 89 229902 (erratum)

DERBIN 02B JETPL 76 409 J. Bernabeu, J. Papavassiliou, J. Vidal

Translated from ZETFP 76 483.

GRIMUS 02 NP B648 376 W. Grimus *et al.*

JOSHIPURA 02B PR D66 113008 A.S. Joshipura, E. Masso, S. Mohanty

LEWIS 02 PR D66 103511 A. Lewis, S. Bridle

LOREDO 02 PR D65 063002 T.J. Loredo, D.Q. Lamb

WANG 02 PR D65 123001 X. Wang, M. Tegmark, M. Zaldarriaga

AUERBACH 01 PR D63 112001 L.B. Auerbach *et al.* (LSND Collab.)

SCHWIENHO... 01 PL B513 23 R. Schwiennhorst *et al.* (DONUT Collab.)

ATHANAS 00 PR D61 052002 M. Athanas *et al.* (CLEO Collab.)

BERNABEU 00 PR D62 113012 J. Bernabeu *et al.*

FUKUGITA 00 PRL 84 1082 M. Fukugita, G.C. Liu, N. Sugiyama

TANIMOTO 00 PL B478 1 N. Tanimoto *et al.*

AYALA 99 PR D59 111901 A. Ayala, J.C. D'Olivo, M. Torres

BEACOM 99 PRL 83 5222 J.F. Beacom, P. Vogel

CROFT 99 PRL 83 1032 R.A.C. Croft, W. Hu, R. Dave

DOLGOV 99 NP B548 385 A.D. Dolgov *et al.*

LOBASHEV 99 PR B460 227 V.M. Lobashev *et al.*

RAFFELT 99 PRPL 320 313 G.C. Raffelt

WEINHEIMER 99 PL B460 219 Ch. Weinheimer *et al.*

ACKERSTAFF 98 EPJ C5 229 K. Ackerstaff *et al.* (OPAL Collab.)

AMMAR 98 PL B431 209 R. Ammar *et al.* (CLEO Collab.)

BARATE 98 EPJ C2 395 R. Barate *et al.* (ALEPH Collab.)

BILLER 98 PRL 80 2992 S.D. Biller *et al.* (WHIPPLe Collab.)

FELDMAN 98 PR D57 3873 G.J. Feldman, R.D. Cousins

LENZ 98 PL B416 50 S. Lenz *et al.*

ABREU 97J ZPHY C74 577 P. Abreu *et al.* (DELPHI Collab.)

ACCIARRI 97Q PL B412 201 M. Acciarri *et al.* (L3 Collab.)

ANASTASSOV 97 PR D55 2559 A. Anastassov *et al.* (CLEO Collab.)

Also PR D58 119303 (erratum)

ELMFORS 97 NP B503 3 P. Elmfors *et al.* (CLEO Collab.)

ESCRIBANO 97 PL B395 369 R. Escribano, E. Masso (BARC, PARIT)

FIELDS 97 ASP 6 169 B.D. Fields, K. Kainulainen, K.A. Olive (NDAM+)

SWAIN 97 PR D55 R1 J. Swain, L. Taylor (NEAS)

ALEXANDER 96M ZPHY C72 231 G. Alexander *et al.* (OPAL Collab.)

ASSAMAGAN 96 PR D53 6065 K.A. Assamagan *et al.* (PSI, ZURI, VILL+)

BAI 96 PR D53 20 J.Z. Bai *et al.* (BES Collab.)

BOTTINO 96 PR D53 6361 A. Bottino *et al.*

DOLGOV 96 PL B383 193 A.D. Dolgov, S. Pastor, J.W.F. Valle (IFIC, VALE)

HANNESTAD 96 PRL 76 2848 S. Hannestad, J. Madsen (AARH)

HANNESTAD 96B PRL 77 5148 (erratum) S. Hannestad, J. Madsen (AARH)

HANNESTAD 96C PR D54 7694 S. Hannestad, J. Madsen (AARH)

SOBIE 96 ZPHY C70 383 R.J. Sobie, R.K. Keeler, I. Lawson (VICT)

BELESEV 95 PL B350 263 A.I. Beleshev *et al.* (INRM, KIAE)

BUSKULIC 95H PL B349 585 D. Buskulic *et al.* (ALEPH Collab.)

CHING 95 IJMP A10 2841 C.R. Ching *et al.* (CST, BEUT, CIAE)

DOLGOV 95 PR D51 4129 A.D. Dolgov, K. Kainulainen, I.Z. Rothstein (MICH+)

HIDDEMANN 95 JP G21 639 K.H. Hiddeemann, H. Daniel, O. Schwentker (MUNT)

KERNAN 95 NP B437 243 P.J. Kernan, L.M. Krauss (CASE)

SIGL 95 PR D51 1499 G. Sigl, M.S. Turner (FNAL, EFI)

STOEFFL 95 PRL 75 3237 W. Stoefl, D.J. Decman (LNL)

VILAIN 95B PL B345 115 P. Vilain *et al.* (CHARM II Collab.)

ASSAMAGAN 94 PL B335 221 K.A. Assamagan *et al.* (PSI, ZURI, VILL+)

BABU 94 PL B321 140 K.S. Babu, T.M. Gould, I.Z. Rothstein (BART+)

DODELSON 94 PR D49 5068 S. Dodelson, G. Gyuk, M.S. Turner (FNAL, CHIC+)

GOULD 94 PL B333 545 T.M. Gould, I.Z. Rothstein (JHU, MICH)

JECKELMANN 94 PL B335 326 B. Jeckelmann, P.F.A. Goudsmit, H.J. Leisi (WABRN+)

KAWASAKI 94 NP B419 105 M. Kawasaki *et al.* (OSU)

PERES 94 PR D50 513 O.L.G. Peres, V. Pleitez, R. Zukanovich Funchal

YASUMI 94 PL B334 229 S. Yasumi *et al.* (KEK, TSUK, KYOT+)

ALLEN 93 PR D47 11 R.C. Allen *et al.* (UCI, LANL, ANL+)

BALEST 93 PR D47 R3671 R. Balest *et al.* (CLEO Collab.)

CINABRO 93 PRL 70 3700 D. Cinabro *et al.* (CLEO Collab.)

DERBIN 93 JETPL 57 768 A.V. Derbin *et al.* (FNP)

Translated from ZETFP 57 755.

DOLGOV 93 PR D47 476 A.D. Dolgov, I.Z. Rothstein (MICH)

ENQVIST 93 PL B301 376 K. Enqvist, H. Uibo (NORD)

SUN 93 CJNP 15 261 H.C. Sun *et al.* (CIAE, CST, BEUT)

WEINHEIMER 93 PL B300 210 C. Weinheimer *et al.* (MANZ)

ALBRECHT 92M PL B292 221 H. Albrecht *et al.* (ARGUS Collab.)

BLUDMAN 92 PR D45 4720 S.A. Bludman (CFPA)

COOPER... 92 PL B280 153 A.M. Cooper-Sarkar *et al.* (BEBEC WA66 Collab.)

DODELSON 92 PRL 68 2572 S. Dodelson, J.A. Frieman, M.S. Turner (FNAL+)

HOLZSCHUH 92B PL B287 381 E. Holzschuh, M. Fritsch, W. Kundig (ZURI)

KAWANO 92 PL B275 487 L.H. Kawano *et al.* (CIT, UCSD, LLL+)

MOURAO 92 PL B285 364 A.M. Mourao, J. Pulido, J.P. Ralston (LISB, LISBT+)

PDG 92 D45 51 K. Hikasa *et al.* (KEK, LBL, BOST+)

VIDYAKIN 92 JETPL 55 206 G.S. Vidyakin *et al.* (KIAE)

Translated from ZETFP 55 212.

ALLEN 91 PR D43 R1 R.C. Allen *et al.* (UCI, LANL, UMD)

DAVIDSON 91 PR D43 2314 S. Davidson, B.A. Campbell, D. Bailey (ALBE+)

DESHPANDE 91 PR D43 943 N.G. Deshpande, K.V.L. Sarma (OREG, TATA)

DORENBOS... 91 ZPHY C51 142 (erratum) J. Dorenbosch *et al.* (CHARM Collab.)

FULLER 91 PR D43 3136 G.M. Fuller, R.A. Malaney (UCSD)

GRANEK 91 IJMP A6 2387 H. GraneK, B.H.J. McKellar (MELB)

See key on page 601

# Lepton Particle Listings

## Neutrino Properties, Number of Neutrino Types

KAWAKAMI	91	PL B256 105	H. Kawakami <i>et al.</i>	(INUS, TOHOK, TINT+)
KOLB	91	PRL 67 533	E.W. Kolb <i>et al.</i>	(FNAL, CHIC)
KRAKAUER	91	PR D44 R6	D.A. Krakauer <i>et al.</i>	(LAMPF E225 Collab.)
LAM	91	PR D44 3345	W.P. Lam, K.W. Ng	(AST)
ROBERTSON	91	PRL 67 957	R.G.H. Robertson <i>et al.</i>	(LASL, LLL)
AHRENS	90	PR D41 3297	L.A. Ahrens <i>et al.</i>	(BNL, BROW, HIRO+)
AVIGNONE	90	PR D41 682	F.T. Avignone, J.I. Collar	(SCUC)
KRAKAUER	90	PL B252 177	D.A. Krakauer <i>et al.</i>	(LAMPF E225 Collab.)
RAFFELT	90	PRL 64 2056	G.G. Raffelt	(MPIM)
WALKER	90	PR D41 689	T.P. Walker	(HARV)
CHUPP	89	PRL 62 505	E.L. Chupp, W.T. Vestrand, C. Reppin	(UNH, MPIM)
DORENBOS...	89	ZPHY C41 567	J. Dorenbosch <i>et al.</i>	(CHARM Collab.)
GRIFOLS	89B	PR D40 3819	J.A. Grifols, E. Masso	(BARC)
KOLB	89	PRL 62 509	E.W. Kolb, M.S. Turner	(CHIC, FNAL)
LOREDO	89	ANYAS 571 601	T.J. Lored, D.Q. Lamb	(CHIC)
RAFFELT	89	PR D39 2066	G.G. Raffelt	(PRIN, UCB)
RAFFELT	89B	APJ 336 61	G. Raffelt, D. Dearborn, J. Silk	(UCB, LLL)
ALBRECHT	88B	PL B202 149	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARBIERI	88	PRL 61 27	R. Barbieri, R.N. Mohapatra	(PISA, UMD)
BORIS	88	PRL 61 245 (erratum)	S.D. Boris <i>et al.</i>	(ITEP, ASCI)
FUKUGITA	88	PRL 60 879	M. Fukugita <i>et al.</i>	(KYOTU, MPIM, UCB)
GROUCH	88	ZPHY C39 553	H. Grouch, R.W. Robinett	(PSU)
RAFFELT	88B	PR D37 549	G.G. Raffelt, D.S.P. Dearborn	(UCB, LLL)
SPERGEL	88	PL B200 366	D.N. Spergel, J.N. Bahcall	(IAS)
VONFEILIT...	88	PL B200 580	F. von Feilitzsch, L. Oberauer	(MUNT)
BARBIELLINI	87	NAT 329 21	G. Barbiellini, G. Cocconi	(CERN)
BORIS	87	PRL 58 2019	S.D. Boris <i>et al.</i>	(ITEP, ASCI)
Also		PRL 61 245 (erratum)	S.D. Boris <i>et al.</i>	(ITEP, ASCI)
BORIS	87B	JETPL 45 333	S.D. Boris <i>et al.</i>	(ITEP)
Translated from ZETFP 45 267				
FUKUGITA	87	PR D36 3817	M. Fukugita, S. Yazaki	(KYOTU, TOKY)
NUSSINOV	87	PR D36 2278	S. Nussinov, Y. Rephaeli	(TELA)
OBERAUER	87	PL B198 113	L.F. Oberauer, F. von Feilitzsch, R.L. Mossbauer	(LLNL)
SPRINGER	87	PR A35 679	P.T. Springer <i>et al.</i>	(KIAE)
KETOV	86	JETPL 44 146	S.N. Ketov <i>et al.</i>	
Translated from ZETFP 44 114				
COWSIK	85	PL 151B 62	R. Cowsik	(TATA)
RAFFELT	85	PR D31 3002	G.G. Raffelt	(MPIM)
BINETRUY	84	PL 134B 174	P. Binetruy, G. Girardi, P. Salati	(LAPP)
FRESE	84	NP B233 167	K. Freese, D.N. Schramm	(CHIC, FNAL)
KYULDJIEV	84	NP B243 387	A.V. Kyuldjiev	(SOFI)
SCHRAMM	84	PL 141B 337	D.N. Schramm, G. Steigman	(FNAL, BART)
VOGEL	84	PR D30 1505	P. Vogel	
ANDERHUB	82	PL 114B 76	H.B. Anderhub <i>et al.</i>	(ETH, SIN)
OLIVE	82	PR D25 213	K.A. Olive, M.S. Turner	(CHIC, UCSB)
BERNSTEIN	81	PL 101B 39	J. Bernstein, G. Feinberg	(STEV, COLU)
FRANK	81	PR D24 2001	J.S. Frank <i>et al.</i>	(LASL, YALE, MIT+)
MORGAN	81	PL 102B 247	J.A. Morgan	(SUSS)
FUJIKAWA	80	PRL 45 963	K. Fujikawa, R. Shrock	(STON)
LUBIMOV	80	PL 94B 266	V.A. Lyubimov <i>et al.</i>	(ITEP)
STECKER	80	PRL 45 1460	F.W. Stecker	(NASA)
COWSIK	79	PR D19 2219	R. Cowsik	(TATA)
GOLDMAN	79	PR D19 2215	T. Goldman, G.J. Stephenson	(LASL)
BEG	78	PR D17 1395	M.A.B. Beg, W.J. Marciano, M. Ruderman	(ROCK+)
BLIETSCHAU	78	NP B133 205	J. Blietschau <i>et al.</i>	(Gargamelle Collab.)
FALK	78	PL 79B 511	S.W. Falk, D.N. Schramm	(CHIC)
BARNES	77	PRL 38 1049	V.E. Barnes <i>et al.</i>	(PURD, ANL)
COWSIK	77	PRL 39 784	R. Cowsik	(MPIM, TATA)
LEE	77C	PR D16 1444	B.W. Lee, R.E. Shrock	(STON)
VYSOTSKY	77	JETPL 26 188	M.I. Vysotsky, A.D. Dolgov, Y.B. Zeldovich	(ITEP)
Translated from ZETFP 26 200				
BELLOTTI	76	LNC 17 553	E. Bellotti <i>et al.</i>	(MILA)
SUTHERLAND	76	PR D13 2700	P. Sutherland <i>et al.</i>	(PENN, COLU, NYU)
SZALAY	76	AA 49 437	A.S. Szalay, G. Marx	(EOTV)
CLARK	74	PR D9 533	A.R. Clark <i>et al.</i>	(LBL)
KIM	74	PR D9 3050	J.E. Kim, V.S. Mathur, S. Okubo	(ROCH)
REINES	74	PRL 32 180	F. Reines, H.W. Sobel, H.S. Gurr	(UCI)
SZALAY	74	APAH 35 8	A.S. Szalay, G. Marx	(EOTV)
COWSIK	72	PRL 29 669	R. Cowsik, J. McClelland	(UCB)
MARX	72	Nu Conf. Budapest	G. Marx, A.S. Szalay	(EOTV)
GERSHTEIN	66	JETPL 4 120	S.S. Gershtein, Y.B. Zeldovich	(KIAM)
Translated from ZETFP 4 189				
BERNSTEIN	63	PR 132 1227	J. Bernstein, M. Ruderman, G. Feinberg	(NYU+)
COWAN	57	PR 107 528	C.L. Cowan, F. Reines	(LANL)

## Number of Neutrino Types

The neutrinos referred to in this section are those of the Standard  $SU(2) \times U(1)$  Electroweak Model possibly extended to allow nonzero neutrino masses. Light neutrinos are those with  $m < m_Z/2$ . The limits are on the number of neutrino mass eigenstates, including  $\nu_1$ ,  $\nu_2$ , and  $\nu_3$ .

## THE NUMBER OF LIGHT NEUTRINO TYPES FROM COLLIDER EXPERIMENTS

Revised March 2008 by D. Karlen (University of Victoria and TRIUMF).

The most precise measurements of the number of light neutrino types,  $N_\nu$ , come from studies of  $Z$  production in  $e^+e^-$  collisions. The invisible partial width,  $\Gamma_{\text{inv}}$ , is determined by subtracting the measured visible partial widths, corresponding to  $Z$  decays into quarks and charged leptons, from the total  $Z$  width. The invisible width is assumed to be due to  $N_\nu$  light neutrino species each contributing the neutrino partial width  $\Gamma_\nu$  as given by the Standard Model. In order to reduce the model dependence, the Standard Model value for the ratio of

the neutrino to charged leptonic partial widths,  $(\Gamma_\nu/\Gamma_\ell)_{\text{SM}} = 1.991 \pm 0.001$ , is used instead of  $(\Gamma_\nu)_{\text{SM}}$  to determine the number of light neutrino types:

$$N_\nu = \frac{\Gamma_{\text{inv}}}{\Gamma_\ell} \left( \frac{\Gamma_\ell}{\Gamma_\nu} \right)_{\text{SM}}. \quad (1)$$

The combined result from the four LEP experiments is  $N_\nu = 2.984 \pm 0.008$  [1].

In the past, when only small samples of  $Z$  decays had been recorded by the LEP experiments and by the Mark II at SLC, the uncertainty in  $N_\nu$  was reduced by using Standard Model fits to the measured hadronic cross sections at several center-of-mass energies near the  $Z$  resonance. Since this method is much more dependent on the Standard Model, the approach described above is favored.

Before the advent of the SLC and LEP, limits on the number of neutrino generations were placed by experiments at lower-energy  $e^+e^-$  colliders by measuring the cross section of the process  $e^+e^- \rightarrow \nu\bar{\nu}\gamma$ . The ASP, CELLO, MAC, MARK J, and VENUS experiments observed a total of 3.9 events above background [2], leading to a 95% CL limit of  $N_\nu < 4.8$ . This process has a much larger cross section at center-of-mass energies near the  $Z$  mass and has been measured at LEP by the ALEPH, DELPHI, L3, and OPAL experiments [3]. These experiments have observed several thousand such events, and the combined result is  $N_\nu = 3.00 \pm 0.08$ . The same process has also been measured by the LEP experiments at much higher center-of-mass energies, between 130 and 208 GeV, in searches for new physics [4]. Combined with the lower energy data, the result is  $N_\nu = 2.92 \pm 0.05$ .

Experiments at  $p\bar{p}$  colliders also placed limits on  $N_\nu$  by determining the total  $Z$  width from the observed ratio of  $W^\pm \rightarrow \ell^\pm \nu$  to  $Z \rightarrow \ell^+ \ell^-$  events [5]. This involved a calculation that assumed Standard Model values for the total  $W$  width and the ratio of  $W$  and  $Z$  leptonic partial widths, and used an estimate of the ratio of  $Z$  to  $W$  production cross sections. Now that the  $Z$  width is very precisely known from the LEP experiments, the approach is now one of those used to determine the  $W$  width.

## References

1. ALEPH, DELPHI, L3, OPAL, and SLD Collaborations, and LEP Electroweak Working Group, and SLD Electroweak Group, and SLD Heavy Flavour Group, Phys. Reports **427**, 257 (2006).
2. VENUS: K. Abe *et al.*, Phys. Lett. **B232**, 431 (1989); ASP: C. Hearty *et al.*, Phys. Rev. **D39**, 3207 (1989); CELLO: H.J. Behrend *et al.*, Phys. Lett. **B215**, 186 (1988); MAC: W.T. Ford *et al.*, Phys. Rev. **D33**, 3472 (1986); MARK J: H. Wu, Ph.D. Thesis, Univ. Hamburg (1986).
3. L3: M. Acciarri *et al.*, Phys. Lett. **B431**, 199 (1998); DELPHI: P. Abreu *et al.*, Z. Phys. **C74**, 577 (1997); OPAL: R. Akers *et al.*, Z. Phys. **C65**, 47 (1995); ALEPH: D. Buskulic *et al.*, Phys. Lett. **B313**, 520 (1993).
4. DELPHI: J. Abdallah *et al.*, Eur. Phys. J. **C38**, 395 (2005); L3: P. Achard *et al.*, Phys. Lett. **B587**, 16 (2004);

# Lepton Particle Listings

## Number of Neutrino Types

- ALEPH: A. Heister *et al.*, Eur. Phys. J. **C28**, 1 (2003);  
OPAL: G. Abbiendi *et al.*, Eur. Phys. J. **C18**, 253 (2000).
5. UA1: C. Albajar *et al.*, Phys. Lett. **B198**, 271 (1987);  
UA2: R. Ansari *et al.*, Phys. Lett. **B186**, 440 (1987).

Number from e <sup>+</sup> e <sup>-</sup> Colliders			
VALUE	DOCUMENT ID	TECN	
2.9840±0.0082	<sup>1</sup> LEP-SLC	06	RVUE
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.00 ±0.05	<sup>2</sup> LEP	92	RVUE
<sup>1</sup> Combined fit from ALEPH, DELPHI, L3 and OPAL Experiments.			
<sup>2</sup> Simultaneous fits to all measured cross section data from all four LEP experiments.			

Number of Light ν Types from Direct Measurement of Invisible Z Width			
In the following, the invisible Z width is obtained from studies of single-photon events from the reaction e <sup>+</sup> e <sup>-</sup> → ν $\overline{\nu}$ γ. All are obtained from LEP runs in the E <sub>CM</sub> <sup>EE</sup> range 88–209 GeV.			
VALUE	DOCUMENT ID	TECN	COMMENT
2.92±0.05 OUR AVERAGE	Error includes scale factor of 1.2.		
2.84±0.10±0.14	ABDALLAH 05b	DLPH	√s = 180–209 GeV
2.98±0.05±0.04	ACHARD 04e	L3	1990–2000 LEP runs
2.86±0.09	HEISTER 03c	ALEP	√s = 189–209 GeV
2.69±0.13±0.11	ABBIENDI,G 00d	OPAL	1998 LEP run
2.89±0.32±0.19	ABREU 97j	DLPH	1993–1994 LEP runs
3.23±0.16±0.10	AKERS 95c	OPAL	1990–1992 LEP runs
2.68±0.20±0.20	BUSKULIC 93l	ALEP	1990–1991 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.84±0.15±0.14	ABREU 00z	DLPH	1997–1998 LEP runs
3.01±0.08	ACCIARRI 99r	L3	1991–1998 LEP runs
3.1 ±0.6 ±0.1	ADAM 96c	DLPH	√s = 130, 136 GeV

Limits from Astrophysics and Cosmology			
Effective Number of Light ν Types			
("Light" means < about 1 MeV). The quoted values correspond to N <sub>eff</sub> , where N <sub>eff</sub> = 3.046 in the Standard Model with N <sub>ν</sub> = 3. See also OLIVE 81. For a review of limits based on Nucleosynthesis, Supernovae, and also on terrestrial experiments, see DENEGR1 90. Also see "Big-Bang Nucleosynthesis" in this Review.			
VALUE	CL%	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.3 ±0.5	95	<sup>1</sup> ADE	14 COSM Planck
3.78+0.31-0.30		<sup>2</sup> COSTANZI	14 COSM
3.29±0.31		<sup>3</sup> HOU	14 COSM
< 3.80		<sup>4</sup> LEISTEDT	14 COSM
< 4.10	95	<sup>5</sup> MORESCO	12 COSM
< 5.79	95	<sup>6</sup> XIA	12 COSM
< 4.08	95	MANGANO	11 COSM BBN
0.9 < N <sub>ν</sub> < 8.2		<sup>7</sup> ICHIKAWA	07 COSM
3 < N <sub>ν</sub> < 7	95	<sup>8</sup> CIRELLI	06 COSM
2.7 < N <sub>ν</sub> < 4.6	95	<sup>9</sup> HANNESTAD	06 COSM
3.6 < N <sub>ν</sub> < 7.4	95	<sup>8</sup> SELJAK	06 COSM
< 4.4		<sup>10</sup> CYBURT	05 COSM
< 3.3		<sup>11</sup> BARGER	03c COSM
1.4 < N <sub>ν</sub> < 6.8		<sup>12</sup> CROTTY	03 COSM
1.9 < N <sub>ν</sub> < 6.6		<sup>12</sup> PIERPAOLI	03 COSM
2 < N <sub>ν</sub> < 4		LISI	99 COSM BBN
< 4.3		OLIVE	99 COSM BBN
< 4.9		COPI	97 Cosmology
< 3.6		HATA	97b High D/H quasar abs.
< 4.0		OLIVE	97 BBN; high <sup>4</sup> He and <sup>7</sup> Li
< 4.7		CARDALL	96b High D/H quasar abs.
< 3.9		FIELDS	96 COSM BBN; high <sup>4</sup> He and <sup>7</sup> Li
< 4.5		KERNAN	96 COSM High D/H quasar abs.
< 3.6		OLIVE	95 BBN; ≥ 3 massless ν
< 3.3		WALKER	91 Cosmology
< 3.4		OLIVE	90 Cosmology
< 4		YANG	84 Cosmology
< 4		YANG	79 Cosmology
< 7		STEIGMAN	77 Cosmology
		PEEBLES	71 Cosmology
<16		<sup>13</sup> SHVARTSMAN	69 Cosmology
		HOYLE	64 Cosmology

- <sup>1</sup> Fit to the number of neutrino degrees of freedom from Planck CMB data along with WMAP polarization, high L, and BAO data.
- <sup>2</sup> Fit to the number of neutrinos degrees of freedom from Planck CMB data along with BAO, shear and cluster data.
- <sup>3</sup> Fit based on the SPT-SZ survey combined with CMB, BAO, and H<sub>0</sub> data.
- <sup>4</sup> Constrains the number of neutrino degrees of freedom (marginalizing over the total mass) from CMB, CMB lensing, BAO, and galaxy clustering data.
- <sup>5</sup> Limit on the number of light neutrino types from observational Hubble parameter data with seven-year WMAP data, SPT, and the most recent estimate of H<sub>0</sub>. Best fit is 3.45 ± 0.65.
- <sup>6</sup> Limit on the number of light neutrino types from the CFHTLS combined with seven-year WMAP data and a prior on the Hubble parameter. Best fit is 4.17<sup>+1.62</sup><sub>-1.26</sub>. Limit is relaxed to 3.98<sup>+2.02</sup><sub>-1.20</sub> when small scales affected by non-linearities are removed.
- <sup>7</sup> Constrains the number of neutrino types from recent CMB and large scale structure data. No priors on other cosmological parameters are used.
- <sup>8</sup> Constrains the number of neutrino types from recent CMB, large scale structure, Lyman-alpha forest, and SN1a data. The slight preference for N<sub>ν</sub> > 3 comes mostly from the Lyman-alpha forest data.
- <sup>9</sup> Constrains the number of neutrino types from recent CMB and large scale structure data. See also HAMANN 07.
- <sup>10</sup> Limit on the number of neutrino types based on <sup>4</sup>He and D/H abundance assuming a baryon density fixed to the WMAP data. Limit relaxes to 4.6 if D/H is not used or to 5.8 if only D/H and the CMB are used. See also CYBURT 01 and CYBURT 03.
- <sup>11</sup> Limit on the number of neutrino types based on combination of WMAP data and big-bang nucleosynthesis. The limit from WMAP data alone is 8.3. See also KNELLER 01. N<sub>ν</sub> ≥ 3 is assumed to compute the limit.
- <sup>12</sup> 95% confidence level range on the number of neutrino flavors from WMAP data combined with other CMB measurements, the 2dFGRS data, and HST data.
- <sup>13</sup> SHVARTSMAN 69 limit inferred from his equations.

Number Coupling with Less Than Full Weak Strength			
VALUE	DOCUMENT ID	TECN	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<20	<sup>1</sup> OLIVE	81c	COSM
<20	<sup>1</sup> STEIGMAN	79	COSM
<sup>1</sup> Limit varies with strength of coupling. See also WALKER 91.			

REFERENCES FOR Limits on Number of Neutrino Types			
ADE	14	AA 571 A16	P.A.R. Ade <i>et al.</i> (Planck Collab.)
COSTANZI	14	JCAP 1410 081	M. Costanzi <i>et al.</i> (TRST, TRST1)
HOU	14	APJ 782 74	Z. Hou <i>et al.</i>
LEISTEDT	14	PRL 113 041301	B. Leistedt, H.V. Peiris, L. Verde
MORESCO	12	JCAP 1207 053	M. Moresco <i>et al.</i>
XIA	12	JCAP 1206 010	J.-Q. Xia <i>et al.</i>
MANGANO	11	PL B701 296	G. Mangano, P. Serpico
HAMANN	07	JCAP 0708 021	J. Hamann <i>et al.</i>
ICHIKAWA	07	JCAP 0705 007	K. Ichikawa, M. Kawasaki, F. Takahashi
CIRELLI	06	JCAP 0612 013	M. Cirelli <i>et al.</i>
HANNESTAD	06	JCAP 0611 016	S. Hannestad, G. Raffelt
LEP-SLC	06	PRPL 427 257	ALEPH, DELPHI, L3, OPAL, SLD and working groups
SELJAK	06	JCAP 0610 014	U. Seljak, A. Slosar, P. McDonald
ABDALLAH	05b	EPL C38 395	J. Abdallah <i>et al.</i> (DELPHI Collab.)
CYBURT	05	ASP 23 313	R.H. Cyburt <i>et al.</i>
ACHARD	04e	PL B587 16	P. Achard <i>et al.</i> (L3 Collab.)
BARGER	03c	PL B566 8	V. Barger <i>et al.</i>
CROTTY	03	PR D67 123005	P. Crotty, J. Lesgourgues, S. Pastor
CYBURT	03	PL B567 227	R.H. Cyburt, B.D. Fields, K.A. Olive
HEISTER	03c	EPJ C28 1	A. Heister <i>et al.</i> (ALEPH Collab.)
PIERPAOLI	03	MNRAS 342 L63	E. Pierpaoli
CYBURT	01	ASP 17 87	R.H. Cyburt, B.D. Fields, K.A. Olive
KNELLER	01	PR D64 123506	J.P. Kneller <i>et al.</i>
ABBIENDI,G	00d	EPJ C18 253	G. Abbiendi <i>et al.</i> (OPAL Collab.)
ABREU	00z	EPJ C17 53	P. Abreu <i>et al.</i> (DELPHI Collab.)
ACCIARRI	99r	PL B470 268	M. Acciarri <i>et al.</i> (L3 Collab.)
LISI	99	PR D59 123520	E. Lisi, S. Sarkar, F.L. Villante
OLIVE	99	ASP 11 403	K.A. Olive, D. Thomas
ABREU	97j	ZPHY C74 577	P. Abreu <i>et al.</i> (DELPHI Collab.)
COPI	97	PR D55 3389	C.J. Copi, D.N. Schramm, M.S. Turner (CHIC)
HATA	97b	PR D55 540	N. Hata <i>et al.</i> (OSU, PENN)
OLIVE	97	ASP 7 27	K.A. Olive, D. Thomas (MINN, FLOR)
ADAM	96c	PL B380 471	W. Adam <i>et al.</i> (DELPHI Collab.)
CARDALL	96b	APJ 472 435	C.Y. Cardall, G.M. Fuller (UCSD)
FIELDS	96	New Ast 1 77	B.D. Fields <i>et al.</i> (NDAM, CERN, MINN+)
KERNAN	96	PR D54 3681	P.S. Kernan, S. Sarkar (CASE, OXFPP)
AKERS	95c	ZPHY C65 47	R. Akers <i>et al.</i> (OPAL Collab.)
OLIVE	95	PL B354 357	K.A. Olive, G. Steigman (MINN, OSU)
BUSKULIC	93l	PL B313 520	D. Buskulić <i>et al.</i> (ALEPH Collab.)
LEP	92	PL B276 247	LEP Collabs. (LEP, ALEPH, DELPHI, L3, OPAL)
WALKER	91	APJ 376 51	T.P. Walker <i>et al.</i> (HSCA, OSU, CHIC+)
DENEGR1	90	RMP 62 1	D. Denegri, B. Sadoulet, M. Spino (CERN, UCB+)
OLIVE	90	PL B236 454	K.A. Olive <i>et al.</i> (MINN, CHIC, OSU+)
YANG	84	APJ 281 493	J. Yang <i>et al.</i> (CHIC, BART)
OLIVE	81	APJ 246 557	K.A. Olive <i>et al.</i> (CHIC, BART)
OLIVE	81c	NP B180 497	K.A. Olive, D.N. Schramm, G. Steigman (EFT+)
STEIGMAN	79	APJ 227 697	G. Steigman, K.A. Olive, D.N. Schramm (BART+)
YANG	77	PL 66B 202	J. Yang <i>et al.</i> (CHIC, YALE, UVA)
STEIGMAN	77	PL 66B 202	G. Steigman, D.N. Schramm, J.E. Gunn (YALE, CHIC+)
PEEBLES	71	Physical Cosmology	P.Z. Peebles (PRIN)
Princeton Univ. Press	(1971)		
SHVARTSMAN	69	JETPL 9 184	V.F. Shvartsman (MOSU)
HOYLE	64	NAT 203 1108	F. Hoyle, R.J. Tayler (CAMB)

**Double- $\beta$  Decay**

OMITTED FROM SUMMARY TABLE

**NEUTRINOLESS DOUBLE- $\beta$  DECAY**

Revised August 2015 by P. Vogel (Caltech) and A. Piepke (University of Alabama).

Observation of neutrinoless double-beta ( $0\nu\beta\beta$ ) decay would signal violation of total lepton number conservation. The process can be mediated by an exchange of a light Majorana neutrino, or by an exchange of other particles. However, the existence of  $0\nu\beta\beta$ -decay requires Majorana neutrino mass, no matter what the actual mechanism is. As long as only a limit on the lifetime is available, limits on the effective Majorana neutrino mass, on the lepton-number violating right-handed current or other possible mechanisms mediating  $0\nu\beta\beta$ -decay can be obtained, independently of the actual mechanism by assuming that one of these “new physics” possibilities dominates. These limits are listed in the next three tables, together with a claimed  $0\nu\beta\beta$ -decay signal reported by part of the Heidelberg-Moscow collaboration. There is tension between that claim and several recent experiments which did not find evidence for  $0\nu\beta\beta$  decay.

In the following we assume that the exchange of light Majorana neutrinos ( $m_{\nu_i} \leq 10$  MeV) contributes dominantly to the decay rate. Besides a dependence on the phase space ( $G^{0\nu}$ ) and the nuclear matrix element ( $M^{0\nu}$ ), the observable  $0\nu\beta\beta$ -decay rate is proportional to the square of the effective Majorana mass  $\langle m_{\beta\beta} \rangle$ ,  $(T_{1/2}^{0\nu})^{-1} = G^{0\nu} \cdot |M^{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$ , with  $\langle m_{\beta\beta} \rangle^2 = |\sum_i U_{ei}^2 m_{\nu_i}|^2$ . The sum contains, in general, complex  $CP$ -phases in  $U_{ei}^2$ , i.e., cancellations may occur. For three neutrino flavors, there are three physical phases for Majorana neutrinos. There is only one phase if neutrinos are Dirac particles. The two additional Majorana phase differences affect only processes to which lepton-number-changing amplitudes contribute. Given the general  $3 \times 3$  mixing matrix for Majorana neutrinos, one can construct other analogous lepton number violating quantities,  $\langle m_{\ell\ell'} \rangle = \sum_i U_{ei} U_{\ell'i} m_{\nu_i}$  ( $\ell$  or  $\ell' \neq e$ ). However, these are currently much less constrained than  $\langle m_{\beta\beta} \rangle$ .

Nuclear structure calculations are needed to deduce  $\langle m_{\beta\beta} \rangle$  from the decay rate. While  $G^{0\nu}$  can be calculated, the computation of  $M^{0\nu}$  is subject to uncertainty. Comparing different nuclear model evaluations indicates a factor  $\sim 2$  to 3 spread in the calculated nuclear matrix elements. In addition, if the effective value of the axial current coupling constant  $g_A$  in nuclei is substantially smaller than its single nucleon value  $g_A = 1.2723 \pm 0.0023$ , the decay rate might be further reduced. The particle physics quantities to be determined are thus nuclear model-dependent, so the half-life measurements are listed first. Where possible, we reference the nuclear matrix elements used in the subsequent analysis. Since rates for the more conventional  $2\nu\beta\beta$  decay serve to calibrate some nuclear models (e.g. QRPA-based calculations), results for this process are also given.

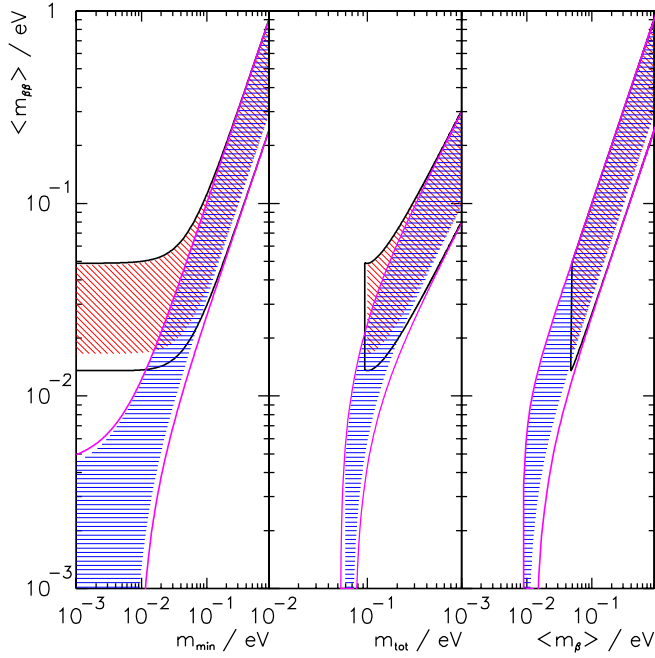
Oscillation experiments utilizing atmospheric-, accelerator-, solar-, and reactor-produced neutrinos and anti-neutrinos yield strong evidence that at least some neutrinos are massive. However, these findings shed no light on the mass hierarchy (i.e., on the sign of  $\Delta m_{31}^2$ ), the absolute neutrino mass values or the properties of neutrinos under CPT-conjugation (Dirac or Majorana).

All confirmed oscillation experiments can be consistently described using three interacting neutrino species with two mass splittings and three mixing angles. Full three flavor analyses such as e.g. [1] yield:  $|\Delta m_{31}^2| = 2.48_{-0.07}^{+0.05} (2.38_{-0.06}^{+0.05}) \times 10^{-3}$  eV<sup>2</sup> and  $\sin^2 \theta_{23} = 0.567_{-0.124}^{+0.032} (0.573_{-0.039}^{+0.025})$  for the parameters observed in atmospheric and accelerator experiments, where the values correspond to the normal (inverted) hierarchies. Observations of solar  $\nu_e$  and reactor  $\bar{\nu}_e$  lead to  $\Delta m_{21}^2 = 7.60_{-0.18}^{+0.19} \times 10^{-5}$  eV<sup>2</sup> and  $\sin^2 \theta_{12} = 0.323 \pm 0.016$ . The investigation of reactor  $\bar{\nu}_e$  at  $\sim 1.5$  km baseline shows that electron type neutrinos couple only weakly to the third mass eigenstate with  $\sin^2 \theta_{13} = 0.0226 \pm 0.0012$  ( $0.0229 \pm 0.0012$ ). (All errors correspond to  $1\sigma$ .)

Based on the 3-neutrino analysis:  $\langle m_{\beta\beta} \rangle^2 = |\cos^2 \theta_{13} \cos^2 \theta_{12} m_1 + e^{i\Delta\alpha_{21}} \cos^2 \theta_{13} \sin^2 \theta_{12} m_2 + e^{i\Delta\alpha_{31}} \sin^2 \theta_{13} m_3|^2$ , with  $\Delta\alpha_{21}, \Delta\alpha_{31}$  denoting the physically relevant Majorana  $CP$ -phase differences (possible Dirac phase  $\delta$  is absorbed in these  $\Delta\alpha$ ). Given the present knowledge of the neutrino oscillation parameters one can derive the relation between the effective Majorana mass and the mass of the lightest neutrino, as illustrated in the left panel of Fig. 1. The three mass hierarchies allowed by the oscillation data: normal ( $m_1 < m_2 < m_3$ ), inverted ( $m_3 < m_1 < m_2$ ), and degenerate ( $m_1 \approx m_2 \approx m_3$ ), result in different projections. The width of the innermost hatched bands reflects the uncertainty introduced by the unknown Majorana and Dirac phases. If the experimental errors of the oscillation parameters are taken into account, then the allowed areas are widened as shown by the outer bands of Fig. 1. Because of the overlap of the different mass scenarios a measurement of  $\langle m_{\beta\beta} \rangle$  in the degenerate or inversely hierarchical ranges would not determine the hierarchy. The middle panel of Fig. 1 depicts the relation of  $\langle m_{\beta\beta} \rangle$  with the summed neutrino mass  $m_{tot} = m_1 + m_2 + m_3$ , constrained by observational cosmology. The oscillation data thus allow to test whether observed values of  $\langle m_{\beta\beta} \rangle$  and  $m_{tot}$  are consistent within the 3 neutrino framework and the light neutrino-exchange dominance assumption. The right hand panel of Fig. 1, finally, shows  $\langle m_{\beta\beta} \rangle$  as a function of the kinematical mass  $\langle m_{\beta} \rangle = [\sum |U_{ei}|^2 m_{\nu_i}^2]^{1/2}$  determined through the analysis of the electron energy distribution in low energy beta decays. The rather large intrinsic width of the  $\beta\beta$ -decay constraint essentially does not allow to positively identify the inverted hierarchy, and thus the sign of  $\Delta m_{31}^2$ , even in combination with these other observables. Naturally, if the value of  $\langle m_{\beta\beta} \rangle \leq 0.01$  eV, but non-zero is ever established then normal hierarchy becomes the only possible scenario.

# Lepton Particle Listings

## Double- $\beta$ Decay



**Figure 1:** The left panel shows the dependence of  $\langle m_{\beta\beta} \rangle$  on the absolute mass of the lightest neutrino  $m_{\min}$ . The middle panel shows  $\langle m_{\beta\beta} \rangle$  as a function of the summed neutrino mass  $m_{\text{tot}}$ , while the right panel depicts  $\langle m_{\beta\beta} \rangle$  as a function of the mass  $\langle m_{\beta} \rangle$ . In all panels the width of the hatched areas is due to the unknown Majorana phases and thus irreducible. The allowed areas given by the solid lines are obtained by taking into account the errors of the oscillation parameters (at 90% confidence level [1]). The two sets of solid lines correspond to the normal (blue) and inverted (red) hierarchies. These sets merge into each other for  $\langle m_{\beta\beta} \rangle \geq 0.1$  eV, which corresponds to the degenerate mass pattern.

It should be noted that systematic uncertainties of the nuclear matrix elements are not folded into the mass projections shown in Fig. 1. Taking this additional uncertainty into account would further widen the allowed areas. The uncertainties in oscillation parameters affect the width of the allowed bands in an asymmetric manner, as shown in Fig. 1. For example, for the degenerate mass pattern ( $\langle m_{\beta\beta} \rangle \geq 0.1$  eV) the upper edge is simply  $\langle m_{\beta\beta} \rangle \sim m$ , where  $m$  is the common mass of the degenerate multiplet, independent of the oscillation parameters, while the lower edge is  $m \cos(2\theta_{12})$ . Similar arguments explain the other features of Fig. 1. The plots in Fig. 1 are based on a 3-neutrino analysis. If it turns out that additional, i.e. sterile light neutrinos exist, the allowed regions would be modified substantially.

If the neutrinoless double-beta decay is observed, it will be possible to fix a range of absolute values of the masses  $m_{\nu_i}$ . Unlike the direct neutrino mass measurements, however, a limit on  $\langle m_{\beta\beta} \rangle$  does not allow one to constrain the individual mass values  $m_{\nu_i}$  even when the mass differences  $\Delta m^2$  are known.

Neutrino oscillation data imply, for the first time, the existence of a lower limit  $\sim 0.014$  eV for the Majorana neutrino mass for the inverted hierarchy mass pattern while  $\langle m_{\beta\beta} \rangle$  could, by fine tuning, vanish in the case of the normal mass hierarchy. Several new double beta searches have been proposed to probe the interesting  $\langle m_{\beta\beta} \rangle$  mass range, with the prospect of full coverage of the inverted mass hierarchy region within the next decade.

The  $0\nu\beta\beta$  decay mechanism discussed so far is not the only way in which the decay can occur. Numerous other possible scenarios have been proposed, however, all of them requiring new physics. It will be a challenging task to decide which mechanism was responsible once  $0\nu\beta\beta$  decay is observed. LHC experiments may reveal corresponding signatures for new physics of lepton number violation. If lepton-number-violating right-handed current weak interactions exist, their strength can be characterized by the phenomenological coupling constants  $\eta$  and  $\lambda$  ( $\eta$  describes the coupling between the right-handed lepton current and left-handed quark current while  $\lambda$  describes the coupling when both currents are right-handed). The  $0\nu\beta\beta$  decay rate then depends on  $\langle \eta \rangle = \eta \sum_i U_{ei} V_{ei}$  and  $\langle \lambda \rangle = \lambda \sum_i U_{ei} V_{ei}$  that vanish for massless or unmixed neutrinos ( $V_{\ell j}$  is a matrix analogous to  $U_{\ell j}$  but describing the mixing with the hypothetical right-handed neutrinos). The observation of the single electron spectra could, in principle, allow to distinguish this mechanism of  $0\nu\beta\beta$  from the light Majorana neutrino exchange driven mode. The limits on  $\langle \eta \rangle$  and  $\langle \lambda \rangle$  are listed in a separate table. The reader is cautioned that a number of earlier experiments did not distinguish between  $\eta$  and  $\lambda$ . In addition, see the section on Majoron searches for additional limits set by these experiments.

## References

1. D.V. Forero, M. Tortola, and J.W.F. Valle, Phys. Rev. D90, 093006 (2014) and private communication with M. Tortola.

## Half-life Measurements and Limits for Double- $\beta$ Decay

In most cases the transitions  $(Z,A) \rightarrow (Z+2,A) + 2e^- + (0 \text{ or } 2) \bar{\nu}_e$  to the  $0^+$  ground state of the final nucleus are listed. However, we also list transitions that increase the nuclear charge ( $2e^+$ ,  $e^+$ /EC and ECEC) and transitions to excited states of the final nuclei ( $0_i^+$ ,  $2^+$ , and  $2_i^+$ ). In the following Listings, only best or comparable limits or lifetimes for each isotope are reported and only those with  $T_{1/2} > 10^{20}$  years that are relevant for particle physics. For  $2\nu$  decay, which is well established, only measured half-lives with the smallest (or comparable) error for each nucleus are reported.

$T_{1/2}(10^{21} \text{ yr})$	CL%	ISOTOPE	TRANSITION	METHOD	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$> 26000$	90	$^{136}\text{Xe}$	$0\nu$	$g.s. \rightarrow 2_1^+$	KamLAND-Zen 1 ASAKURA 16
$> 26000$	90	$^{136}\text{Xe}$	$0\nu$	$g.s. \rightarrow 2_2^+$	KamLAND-Zen 2 ASAKURA 16
$> 24000$	90	$^{136}\text{Xe}$	$0\nu$	$g.s. \rightarrow 0_1^+$	KamLAND-Zen 3 ASAKURA 16
$1.926 \pm 0.094$		$^{76}\text{Ge}$	$2\nu$	$g.s. \rightarrow g.s.$	GERDA 4 AGOSTINI 15A
$> 4000$	90	$^{130}\text{Te}$	$0\nu$	$g.s. \rightarrow g.s.$	CUORE 5 ALFONSO 15
$(6.93 \pm 0.04) \times 10^{-3}$		$^{100}\text{Mo}$	$2\nu$		NEMO-3 6 ARNOLD 15
$> 1100$	90	$^{100}\text{Mo}$	$0\nu$		NEMO-3 7 ARNOLD 15
$2.165 \pm 0.016 \pm 0.059$		$^{136}\text{Xe}$	$2\nu$	$g.s. \rightarrow g.s.$	EXO-200 8 ALBERT 14
$> 11000$	90	$^{136}\text{Xe}$	$0\nu$	$g.s. \rightarrow g.s.$	EXO-200 9 ALBERT 14B
$> 1100$	90	$^{100}\text{Mo}$	$0\nu$	$\langle m \rangle$ -driven	NEMO-3 10 ARNOLD 14



See key on page 601

# Lepton Particle Listings

## Double- $\beta$ Decay

> 600	90	<sup>100</sup> Mo	0 $\nu$	$\langle\lambda\rangle$ -driven	NEMO-3	11	ARNOLD	14	1 ASAKURA 16 use the KamLAND-Zen liquid scintillator calorimeter ( <sup>136</sup> Xe 89.5 kg yr) to place a limit on the 0 $\nu\beta\beta$ -decay into the first excited state of the daughter nuclide.
> 1000	90	<sup>100</sup> Mo	0 $\nu$	$\langle\eta\rangle$ -driven	NEMO-3	12	ARNOLD	14	2 ASAKURA 16 use the KamLAND-Zen liquid scintillator calorimeter ( <sup>136</sup> Xe 89.5 kg yr) to place a limit on the 0 $\nu\beta\beta$ -decay into the second excited state of the daughter nuclide.
0.107 $\pm$ 0.046 -0.026		<sup>150</sup> Nd	0 $\nu$ +2 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	$\gamma$ in Ge det.	13	KIDD	14	3 ASAKURA 16 use the KamLAND-Zen liquid scintillator calorimeter ( <sup>136</sup> Xe 89.5 kg yr) to place a limit on the 0 $\nu\beta\beta$ -decay into the third excited state of the daughter nuclide.
> 21000	90	<sup>76</sup> Ge	0 $\nu$	g.s. $\rightarrow$ g.s.	GERDA	14	AGOSTINI	13A	4 AGOSTINI 15A use 17.9 kg yr exposure of the GERDA calorimeter to derive an improved measurement of the 2 $\nu\beta\beta$ decay half life of <sup>76</sup> Ge.
> 0.13	90	<sup>96</sup> Ru	0 $\nu$ +2 $\nu$	2 $\beta^+$ , g.s.	Ge counting	15	BELLI	13A	5 ALFONSO 15 use the combined exposure of the high resolution CUORICINO (19.75 kg yr) and CUORE-0 (9.8 kg yr) bolometers to construct a Bayesian limit on the 0 $\nu\beta\beta$ decay half life of <sup>130</sup> Te.
> 19000	90	<sup>136</sup> Xe	0 $\nu$	g.s. $\rightarrow$ g.s.	KamLAND-Zen	16	GANDO	13A	6 ARNOLD 15 use the NEMO-3 tracking calorimeter with 34.3 kg yr exposure to determine the 2 $\nu\beta\beta$ -half life of <sup>100</sup> Mo. Supersedes ARNOLD 05A and ARNOLD 04.
9.2 $\pm$ 5.5 -2.6 $\pm$ 1.3		<sup>78</sup> Kr	2 $\nu$ 2K	g.s. $\rightarrow$ g.s.	BAKSAN	17	GAVRILYAK	13	7 ARNOLD 15 use the NEMO-3 tracking calorimeter with 34.3 kg yr exposure to determine the limit of 0 $\nu\beta\beta$ -half life of <sup>100</sup> Mo. Supersedes ARNOLD 2005A and BARABASH 11A.
> 5.4	90	<sup>78</sup> Kr	0 $\nu$ 2K	g.s. $\rightarrow$ 2 $^+$	BAKSAN	18	GAVRILYAK	13	8 ALBERT 14 use the EXO-200 tracking detector for a re-measurement of the 2 $\nu\beta\beta$ -half life of <sup>136</sup> Xe. A nuclear matrix element of 0.0218 $\pm$ 0.0003 MeV <sup>-1</sup> is derived from this data. Supersedes ACKERMAN 11.
> 940	90	<sup>130</sup> Te	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	CUORICINO	19	ANDREOTTI	12	9 ALBERT 14B use 100 kg yr of exposure of the EXO-200 tracking calorimeter to place a lower limit on the 0 $\nu\beta\beta$ -half life of <sup>136</sup> Xe. Supersedes AUGER 12.
> 1.0	90	<sup>106</sup> Cd	0 $\nu$	ECEC, g.s.	<sup>106</sup> CdWO <sub>4</sub> scint.	20	BELLI	12A	10 ARNOLD 14 use 34.7 kg yr of exposure of the NEMO-3 tracking calorimeter to derive a limit on the $\langle m \rangle$ -driven (light neutrino mass) 0 $\nu\beta\beta$ -half life of <sup>100</sup> Mo. Supersedes BARABASH 11A.
> 2.2	90	<sup>106</sup> Cd	0 $\nu$	$\beta^+$ +EC, g.s.	<sup>106</sup> CdWO <sub>4</sub> scint.	21	BELLI	12A	11 ARNOLD 14 use 34.7 kg yr of exposure of the NEMO-3 tracking calorimeter to derive a limit on the $\langle\lambda\rangle$ -driven (right handed quark and lepton currents) 0 $\nu\beta\beta$ -half life of <sup>100</sup> Mo.
> 1.2	90	<sup>106</sup> Cd	0 $\nu$	2 $\beta^+$ , g.s.	<sup>106</sup> CdWO <sub>4</sub> scint.	22	BELLI	12A	12 ARNOLD 14 use 34.7 kg yr of exposure of the NEMO-3 tracking calorimeter to derive a limit on the $\langle\eta\rangle$ -driven (right handed quark current) 0 $\nu\beta\beta$ -half life of <sup>100</sup> Mo.
2.38 $\pm$ 0.02 $\pm$ 0.14		<sup>136</sup> Xe	2 $\nu$	g.s. $\rightarrow$ g.s.	KamLAND-Zen	23	GANDO	12A	13 KIDD 14 utilize two underground Ge detectors to determine the inclusive double beta decay rate to the first excited 0 $_1^+$ state using $\gamma\gamma$ coincidences.
0.7 $\pm$ 0.09 $\pm$ 0.11		<sup>130</sup> Te	2 $\nu$		NEMO-3	24	ARNOLD	11	14 AGOSTINI 13A use 21.6 kg yr of data, collected with GERDA detector array, to place a lower limit on the 0 $\nu\beta\beta$ -half life of <sup>76</sup> Ge. This result is in tension with the evidence for 0 $\nu\beta\beta$ -decay reported in KLAPDOR-KLEINGROTHAUS 06A. This half-life limit exceeds the limit reported in KLAPDOR-KLEINGROTHAUS 01.
> 130	90	<sup>130</sup> Te	0 $\nu$		NEMO-3	25	ARNOLD	11	15 BELLI 13A use an underground Ge detector to search for the 2 $\beta^+$ -decay of <sup>96</sup> Ru via the intensity of the annihilation peak. This method cannot distinguish two from zero neutrino decay.
> 1.3	90	<sup>112</sup> Sn	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	$\gamma$ Ge det.	26	BARABASH	11	16 GANDO 13A use the KamLAND detector to search for 0 $\nu\beta\beta$ -decay of <sup>136</sup> Xe based on an exposure of 89.5 kg yr. This result is in tension with the evidence of 0 $\nu\beta\beta$ reported in KLAPDOR-KLEINGROTHAUS 06A and earlier references to that work. Supersedes GANDO 12A and is more sensitive than BERNABEI 02D.
> 0.69	90	<sup>112</sup> Sn	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_2^+$	$\gamma$ Ge det.	27	BARABASH	11	17 GAVRILYAK 13 use a proportional counter filled with Kr gas to search for the 2 $\nu$ 2K decay of <sup>78</sup> Kr. Data with the enriched and depleted Kr were used to determine signal and background. A 2.5 $\sigma$ excess of events obtained with the enriched sample is interpreted as an indication for the presence of this decay.
> 1.3	90	<sup>112</sup> Sn	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	$\gamma$ Ge det.	28	BARABASH	11	18 GAVRILYAK 13 use a proportional counter filled with Kr gas to search for the 0 $\nu$ 2K decay of <sup>78</sup> Kr into 2828 keV excited state of <sup>78</sup> Se. This transition could be subject to resonant rate enhancement. Data obtained with the enriched and depleted Kr were used to determine signal and background.
> 1.06	90	<sup>112</sup> Sn	0 $\nu$		$\gamma$ Ge det.	29	BARABASH	11	19 ANDREOTTI 12 use high resolution TeO <sub>2</sub> bolometric calorimeter to search for the 0 $\nu\beta\beta$ decay of <sup>130</sup> Te leading to the excited 0 $_1^+$ state at 1793.5 keV.
(2.8 $\pm$ 0.1 $\pm$ 0.3)E-2		<sup>116</sup> Cd	2 $\nu$		NEMO-3	30	BARABASH	11A	20 BELLI 12A use <sup>106</sup> CdWO <sub>4</sub> 215 g crystal scintillator to search for various $\beta\beta$ decay modes. The limit for the ECEC mode is derived from the fit to the background spectrum in the 1.8–3.0 MeV energy interval in the run of 6590 hours. The same analysis provides several limits ( $\sim$ 2–5 $\times$ 10 <sup>20</sup> years) for the ECEC mode leading to the excited 0 $^+$ and 2 $^+$ states. Also a similar size limits for the possible resonance process populating states at 2718 keV, 2741 keV, and 2748 keV were obtained.
(4.4 $\pm$ 0.5 -0.4 $\pm$ 0.4)E-2		<sup>48</sup> Ca	2 $\nu$		NEMO-3	31,32	BARABASH	11A	21 BELLI 12A use <sup>106</sup> CdWO <sub>4</sub> 215 g crystal scintillator to search for various $\beta\beta$ decay modes. The limit for the EC $\beta^+$ mode is derived from the fit to the background spectrum in the 2.0–3.0 MeV energy interval in the run of 6590 hours. The same analysis provides several limits ( $\sim$ 0.5–1.3 $\times$ 10 <sup>21</sup> years) for the EC $\beta^+$ mode leading to the excited 0 $^+$ and 2 $^+$ states.
(69 $\pm$ 9 $\pm$ 10)E-2		<sup>130</sup> Te	2 $\nu$		NEMO-3	32,33	BARABASH	11A	22 BELLI 12A use <sup>106</sup> CdWO <sub>4</sub> 215 g crystal scintillator to search for various $\beta\beta$ decay modes. The limit for the $\beta^+\beta^+$ mode is derived from the fit to the background spectrum in the 0.76–2.8 MeV energy interval in the run of 6590 hours. The same analysis provides the limit (1.2 $\times$ 10 <sup>21</sup> years) for the $\beta^+\beta^+$ mode leading to the first excited 2 $^+$ state.
> 360	90	<sup>82</sup> Se	0 $\nu$		NEMO-3	32,34	BARABASH	11A	23 GANDO 12A use a modification of the existing KamLAND detector. The $\beta\beta$ decay source/detector is 13 tons of enriched <sup>136</sup> Xe-loaded scintillator contained in an inner balloon. The 2 $\nu\beta\beta$ decay rate is derived from the fit to the spectrum between 0.5 and 4.8 MeV. This result is in agreement with ACKERMAN 11.
> 100	90	<sup>130</sup> Te	0 $\nu$		NEMO-3	32,35	BARABASH	11A	24 ARNOLD 11 use enriched <sup>130</sup> Te in the NEMO-3 detector to measure the 2 $\nu$ $\beta\beta$ decay rate. This result is in agreement with, but more accurate than ARNABOLDI 03.
> 16	90	<sup>116</sup> Cd	0 $\nu$		NEMO-3	32,36	BARABASH	11A	25 ARNOLD 11 use the NEMO-3 detector to obtain a limit for the 0 $\nu$ $\beta\beta$ decay. This result is less significant than ARNABOLDI 05.
> 0.32	90	<sup>64</sup> Zn	0 $\nu$	ECEC, g.s.	ZnWO <sub>4</sub> scint.	37	BELLI	11D	26 BARABASH 11 use 100 g of enriched <sup>112</sup> Sn to determine a limit for the ECEC 0 $\nu\beta\beta$ decay to the 0 $_1^+$ state of <sup>112</sup> Cd by searching for the de-excitation $\gamma$ with a Ge detector. This decay mode is a candidate for resonant rate enhancement.
> 0.85	90	<sup>64</sup> Zn	0 $\nu$	$\beta^+$ +EC, g.s.	ZnWO <sub>4</sub> scint.	37	BELLI	11D	27 BARABASH 11 use 100 g of enriched <sup>112</sup> Sn to determine a limit for the ECEC 0 $\nu\beta\beta$ decay to the 0 $_2^+$ state of <sup>112</sup> Cd by searching for the de-excitation $\gamma$ with a Ge detector.
> 0.11	90	<sup>106</sup> Cd	0 $\nu$	0 $^+$ $\rightarrow$ 4 $^+$	TGV2 det.	38	RUHADZE	11	28 BARABASH 11 use 100 g of enriched <sup>112</sup> Sn to determine a limit for the ECEC 0 $\nu\beta\beta$ decay to the 0 $_1^+$ state of <sup>112</sup> Cd by searching for the de-excitation $\gamma$ with a Ge detector.
(2.35 $\pm$ 0.14 $\pm$ 0.16)E-2		<sup>96</sup> Zr	2 $\nu$		NEMO-3	39	ARGYRIADES	10	29 BARABASH 11 use 100 g of enriched <sup>112</sup> Sn to determine a limit for the ECEC 0 $\nu\beta\beta$ decay to the ground state of <sup>112</sup> Cd by searching for the de-excitation $\gamma$ with a Ge detector.
> 9.2	90	<sup>96</sup> Zr	0 $\nu$		NEMO-3	40	ARGYRIADES	10	
> 0.22	90	<sup>96</sup> Zr	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	NEMO-3	41	ARGYRIADES	10	
0.69 $\pm$ 0.10 -0.08 $\pm$ 0.07		<sup>100</sup> Mo	2 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	Ge coinc.	42	BELLI	10	
> 18.0	90	<sup>150</sup> Nd	0 $\nu$		NEMO-3	43	ARGYRIADES	09	
(9.11 $\pm$ 0.25 -0.22 $\pm$ 0.63)E-3		<sup>150</sup> Nd	2 $\nu$		NEMO-3	44	ARGYRIADES	09	
> 0.43	90	<sup>64</sup> Zn	0 $\nu$	$\beta^+$ +EC	ZnWO <sub>4</sub> scint.	45	BELLI	09A	
> 0.11	90	<sup>64</sup> Zn	0 $\nu$	ECEC	ZnWO <sub>4</sub> scint.	46	BELLI	09A	
0.55 $\pm$ 0.12 -0.09		<sup>100</sup> Mo	2 $\nu$ +0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	Ge coincidence	47	KIDD	09	
> 0.22	90	<sup>64</sup> Zn	0 $\nu$		ZnWO <sub>4</sub> scint.	48	BELLI	08	
> 1.1	90	<sup>114</sup> Cd	0 $\nu$	2 $\beta$	CdWO <sub>4</sub> scint.	49	BELLI	08B	
> 58	90	<sup>48</sup> Ca	0 $\nu$		CaF <sub>2</sub> scint.	50	UMEHARA	08	
0.57 $\pm$ 0.13 -0.09 $\pm$ 0.08		<sup>100</sup> Mo	2 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	NEMO-3	51	ARNOLD	07	
> 89	90	<sup>100</sup> Mo	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	NEMO-3	52	ARNOLD	07	
> 160	90	<sup>100</sup> Mo	0 $\nu$	0 $^+$ $\rightarrow$ 2 $^+$	NEMO-3	53	ARNOLD	07	
22300 $\pm$ 4400 -3100		<sup>76</sup> Ge	0 $\nu$		Enriched HPGe	54	KLAPDOR-K...	06A	
> 1800	90	<sup>130</sup> Te	0 $\nu$		Cryog. det.	55	ARNABOLDI	05	
> 100	90	<sup>82</sup> Se	0 $\nu$		NEMO-3	56	ARNOLD	05A	
(9.6 $\pm$ 0.3 $\pm$ 1.0)E-2		<sup>82</sup> Se	2 $\nu$		NEMO-3	57	ARNOLD	05A	
> 140	90	<sup>82</sup> Se	0 $\nu$		NEMO-3	58	ARNOLD	04	
0.14 $\pm$ 0.04 -0.02 $\pm$ 0.03		<sup>150</sup> Nd	0 $\nu$ +2 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	$\gamma$ in Ge det.	59	BARABASH	04	
> 31	90	<sup>130</sup> Te	0 $\nu$	0 $^+$ $\rightarrow$ 2 $^+$	Cryog. det.	60	ARNABOLDI	03	
> 110	90	<sup>128</sup> Te	0 $\nu$		Cryog. det.	61	ARNABOLDI	03	
(0.029 $\pm$ 0.004, -0.003)		<sup>116</sup> Cd	2 $\nu$		<sup>116</sup> CdWO <sub>4</sub> scint.	62	DANEVICH	03	
> 170	90	<sup>116</sup> Cd	0 $\nu$		<sup>116</sup> CdWO <sub>4</sub> scint.	63	DANEVICH	03	
> 29	90	<sup>116</sup> Cd	0 $\nu$	0 $^+$ $\rightarrow$ 2 $^+$	<sup>116</sup> CdWO <sub>4</sub> scint.	64	DANEVICH	03	
> 14	90	<sup>116</sup> Cd	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_1^+$	<sup>116</sup> CdWO <sub>4</sub> scint.	65	DANEVICH	03	
> 6	90	<sup>116</sup> Cd	0 $\nu$	0 $^+$ $\rightarrow$ 0 $_2^+$	<sup>116</sup> CdWO <sub>4</sub> scint.	66	DANEVICH	03	
> 1.1	90	<sup>186</sup> W	0 $\nu$		CdWO <sub>4</sub> scint.	67	DANEVICH	03	
> 1.1	90	<sup>186</sup> W	0 $\nu$	0 $^+$ $\rightarrow$ 2 $^+$	CdWO <sub>4</sub> scint.	68	DANEVICH	03	
>15700	90	<sup>76</sup> Ge	0 $\nu$		Enriched HPGe	69	AALSETH	02B	
> 58	90	<sup>134</sup> Xe	0 $\nu$		Liquid Xe Scint.	70	BERNABEI	02D	
> 1.3	90	<sup>160</sup> Gd	0 $\nu$		Gd <sub>2</sub> SiO <sub>5</sub> :Ce	71	DANEVICH	01	
> 1.3	90	<sup>160</sup> Gd	0 $\nu$	0 $^+$ $\rightarrow$ 2 $^+$	Gd <sub>2</sub> SiO <sub>5</sub> :Ce	72	DANEVICH	01	
> 19000	90	<sup>76</sup> Ge	0 $\nu$		Enriched HPGe	73	KLAPDOR-K...	01	
(9.4 $\pm$ 3.2)E-3		<sup>96</sup> Zr	0 $\nu$ +2 $\nu$		Geochem	74	WIESER	01	
0.042 $\pm$ 0.033 -0.013		<sup>48</sup> Ca	2 $\nu$		Ge spectrometer	75	BRUDANIN	00	
0.021 $\pm$ 0.008 -0.004 $\pm$ 0.002		<sup>96</sup> Zr	2 $\nu$		NEMO-2	76	ARNOLD	99	
> 2.8	90	<sup>82</sup> Se	0 $\nu$	0 $^+$ $\rightarrow$ 2 $^+$	NEMO-2	77	ARNOLD	98	
(6.75 $\pm$ 0.37 -0.42 $\pm$ 0.68)E-3		<sup>150</sup> Nd	2 $\nu$		TPC	78	DESILVA	97	
0.043 $\pm$ 0.024 -0.011 $\pm$ 0.014		<sup>48</sup> Ca	2 $\nu$		TPC	79	BALYSH	96	
0.026 $\pm$ 0.009 -0.005		<sup>116</sup> Cd	2 $\nu$	0 $^+$ $\rightarrow$ 0 $^+$	ELEGANT IV	80	EJIRI	95	
7200 $\pm$ 400		<sup>128</sup> Te	0 $\nu$ +2 $\nu$		Geochem	81	BERNATOW...	92	
2.0 $\pm$ 0.6		<sup>238</sup> U	0 $\nu$ +2 $\nu$		Radiochem	82	TURKOVICH	91	
1800 $\pm$ 700		<sup>128</sup> Te	0 $\nu$ +2 $\nu$		Geochem.	82	LIN	88B	

# Lepton Particle Listings

## Double-β Decay

<sup>30</sup> Supersedes DANEVICH 03 and ARNOLD 96.

<sup>31</sup> Supersedes BRUDANIN 00 and BALYSH 96.

<sup>32</sup> BARABASH 11A use the NEMO-3 detector to measure  $2\nu\beta\beta$  rates and place limits on  $0\nu\beta\beta$  half lives for various nuclides.

<sup>33</sup> Supersedes ARNABOLDI 03.

<sup>34</sup> Supersedes ARNOLD 05A, ARNOLD 04, ARNOLD 98, and ELLIOTT 92.

<sup>35</sup> Less restrictive than ARNABOLDI 08.

<sup>36</sup> Less restrictive than DANEVICH 03.

<sup>37</sup> BELLI 11D use  $\text{ZnWO}_4$  scintillator calorimeters to search for various  $\beta\beta$  decay modes of  $^{64}\text{Zn}$ ,  $^{70}\text{Zn}$ ,  $^{180}\text{W}$ , and  $^{186}\text{W}$ .

<sup>38</sup> RUKHADZE 11 uses 13.6 g of enriched  $^{106}\text{Cd}$  to search for the neutrinoless ECEC decay into an excited state of  $^{106}\text{Pd}$  and its characteristic  $\gamma$ -radiation using the TGV2 detector. This decay mode is a candidate for resonant rate enhancement, however, hindered by the large spin difference.

<sup>39</sup> ARGYRIADES 10 use  $9.4 \pm 0.2$  g of  $^{96}\text{Zr}$  in NEMO-3 detector and identify its  $2\nu\beta\beta$  decay. The result is in agreement and supersedes ARNOLD 99.

<sup>40</sup> ARGYRIADES 10 use  $9.4 \pm 0.2$  g of  $^{96}\text{Zr}$  in NEMO-3 detector and obtain a limit of the  $0\nu\beta\beta$  decay. The result is in agreement and supersedes ARNOLD 99.

<sup>41</sup> ARGYRIADES 10 use  $9.4 \pm 0.2$  g of  $^{96}\text{Zr}$  in NEMO-3 detector and obtain a limit of the  $0\nu\beta\beta$  decay into the first excited  $0_1^+$  state in  $^{96}\text{Mo}$ .

<sup>42</sup> BELLI 10 use enriched  $^{100}\text{Mo}$  with 4 HP Ge detectors to record the 590.8 and 539.5 keV  $\gamma$  rays from the decay of the  $0_1^+$  state in  $^{100}\text{Ru}$  both in singles and coincidences. This result confirms the measurement of KIDD 09 and ARNOLD 07 and supersedes them.

<sup>43</sup> ARGYRIADES 09 use the NEMO-3 tracking calorimeter containing 36.5 g of  $^{150}\text{Nd}$ , a total exposure of 924.7 days, to derive a limit for the  $0\nu\beta\beta$  half-life. Supersedes DESILVA 97.

<sup>44</sup> ARGYRIADES 09 use the NEMO-3 tracking calorimeter containing 36.5 g of  $^{150}\text{Nd}$ , a total exposure of 924.7 days, to determine the value of the  $2\nu\beta\beta$  half-life. This result is in marginal agreement, but has somewhat smaller error bars, than DESILVA 97.

<sup>45</sup> BELLI 09A use  $\text{ZnWO}_4$  scintillating crystals to search for various modes of  $\beta\beta$  decay. This work improves the limits for different modes of  $^{64}\text{Zn}$  decay into the ground state of  $^{64}\text{Ni}$ , in this case for the  $0\nu\beta^+\text{EC}$  mode. Supersedes BELLI 08.

<sup>46</sup> BELLI 09A use  $\text{ZnWO}_4$  scintillating crystals to search for various modes of  $\beta\beta$  decay. This work improves the limits for different modes of  $^{64}\text{Zn}$  decay into the ground state of  $^{64}\text{Ni}$ , in this case for the  $0\nu\beta\beta$  ECEC mode. Supersedes BELLI 08.

<sup>47</sup> KIDD 09 combine past and new data with an improved coincidence detection efficiency determination. The result agrees with ARNOLD 95. Supersedes DEBRAECKELEER 01 and BARABASH 95.

<sup>48</sup> BELLI 08 use  $\text{ZnWO}_4$  scintillation calorimeter to search for neutrinoless  $\beta^+$  plus electron capture decay of  $^{64}\text{Zn}$ . The half-life limit for the  $2\nu\beta\beta$  mode is  $2.1 \times 10^{20}$  years.

<sup>49</sup> BELLI 08B use  $\text{CdWO}_4$  scintillation calorimeter to search for  $0\nu\beta\beta$  decay of  $^{114}\text{Cd}$ .

<sup>50</sup> UMEHARA 08 use  $\text{CaF}_2$  scintillation calorimeter to search for double beta decay of  $^{48}\text{Ca}$ . Limit is significantly more stringent than quoted sensitivity:  $18 \times 10^{21}$  years.

<sup>51</sup> First exclusive measurement of  $2\nu$ -decay to the first excited  $0_1^+$ -state of daughter nucleus. ARNOLD 07 use the NEMO-3 tracking calorimeter to detect all particles emitted in decay. Result agrees with the inclusive ( $0\nu + 2\nu$ ) measurement of DEBRAECKELEER 01.

<sup>52</sup> Limit on  $0\nu$ -decay to the first excited  $0_1^+$ -state of daughter nucleus using NEMO-3 tracking calorimeter. Supersedes DASSIE 95.

<sup>53</sup> Limit on  $0\nu$ -decay to the first excited  $2^+$ -state of daughter nucleus using NEMO-3 tracking calorimeter.

<sup>54</sup> KLAPDOR-KLEINGROTHAUS 06A present re-analysis of data originally published in KLAPDOR-KLEINGROTHAUS 04A. Modified pulse shape analysis leads the authors to claim improved  $6\sigma$  statistical evidence for observation of  $0\nu$ -decay, compared to  $4.2\sigma$  in KLAPDOR-KLEINGROTHAUS 04A. Analysis of the systematic uncertainty is not presented. This re-analysis is disputed in AGOSTINI 13A and SCHWINGENHEUER 13.

<sup>55</sup> Supersedes ARNABOLDI 04. Bolometric  $\text{TeO}_2$  detector array CUORICINO is used for high resolution search for  $0\nu\beta\beta$  decay. The half-life limit is derived from 3.09 kg yr  $^{130}\text{Te}$  exposure.

<sup>56</sup> NEMO-3 tracking calorimeter is used in ARNOLD 05A to place limit on  $0\nu\beta\beta$  half-life of  $^{82}\text{Se}$ . Detector contains 0.93 kg of enriched  $^{82}\text{Se}$ . Supersedes ARNOLD 04.

<sup>57</sup> ARNOLD 05A use the NEMO-3 tracking detector to determine the  $2\nu\beta\beta$  half-life of  $^{82}\text{Se}$  with high statistics and low background (389 days of data taking). Supersedes ARNOLD 04.

<sup>58</sup> ARNOLD 04 use the NEMO-3 tracking detector to determine the limit for  $0\nu\beta\beta$  half-life of  $^{82}\text{Se}$ . This represents an improvement, by a factor of  $\sim 10$ , when compared with ELLIOTT 92. It supersedes the limit of ARNOLD 98 for this decay using NEMO-2.

<sup>59</sup> BARABASH 04 perform an inclusive measurement of the  $\beta\beta$  decay of  $^{150}\text{Nd}$  into the first excited ( $0_1^+$ ) state of  $^{150}\text{Sm}$ . Gamma radiation emitted in decay of the excited state is detected.

<sup>60</sup> Decay into first excited state of daughter nucleus.

<sup>61</sup> Supersedes ALESSANDRELLO 00. Array of  $\text{TeO}_2$  crystals in high resolution cryogenic calorimeter. Some enriched in  $^{128}\text{Te}$ . Ground state to ground state decay.

<sup>62</sup> Calorimetric measurement of  $2\nu\beta\beta$  ground state decay of  $^{116}\text{Cd}$  using enriched  $\text{CdWO}_4$  scintillators. Agrees with EJIRI 95 and ARNOLD 96. Supersedes DANEVICH 00.

<sup>63</sup> Limit on  $0\nu\beta\beta$  decay of  $^{116}\text{Cd}$  using enriched  $\text{CdWO}_4$  scintillators. Supersedes DANEVICH 00.

<sup>64</sup> Limit on  $0\nu\beta\beta$  decay of  $^{116}\text{Cd}$  into first excited  $2^+$  state of daughter nucleus using enriched  $\text{CdWO}_4$  scintillators. Supersedes DANEVICH 00.

<sup>65</sup> Limit on  $0\nu\beta\beta$  decay of  $^{116}\text{Cd}$  into first excited  $0^+$  state of daughter nucleus using enriched  $\text{CdWO}_4$  scintillators. Supersedes DANEVICH 00.

<sup>66</sup> Limit on  $0\nu\beta\beta$  decay of  $^{116}\text{Cd}$  into second excited  $0^+$  state of daughter nucleus using enriched  $\text{CdWO}_4$  scintillators. Supersedes DANEVICH 00.

<sup>67</sup> Limit on the  $0\nu\beta\beta$  ground state decay of  $^{186}\text{W}$  using enriched  $\text{CdWO}_4$  scintillators.

<sup>68</sup> Limit on the  $0\nu\beta\beta$  decay of  $^{186}\text{W}$  to the first excited  $2^+$  state of the daughter nucleus using enriched  $\text{CdWO}_4$  scintillators.

<sup>69</sup> AALSETH 02b limit is based on 117 mol-yr of data using enriched Ge detectors. Background reduction by means of pulse shape analysis is applied to part of the data set. Reported limit is slightly less restrictive than that in KLAPDOR-KLEINGROTHAUS 01 However, it excludes part of the allowed half-life range reported in KLAPDOR-KLEINGROTHAUS 01b for the same nuclide. The analysis has been criticized in KLAPDOR-KLEINGROTHAUS 04b. The criticism was addressed and disputed in AALSETH 04.

<sup>70</sup> BERNABEI 02n report a limit for the  $0\nu, 0^+ \rightarrow 0^+$  decay of  $^{134}\text{Xe}$ , present in the source at 17%, by considering the maximum number of events for this mode compatible with the fitted smooth background.

<sup>71</sup> DANEVICH 01 place limit on  $0\nu\beta\beta$  decay of  $^{160}\text{Gd}$  using  $\text{Gd}_2\text{SiO}_5\text{:Ce}$  crystal scintillators. The limit is more stringent than KOBAYASHI 95.

<sup>72</sup> DANEVICH 01 place limits on  $0\nu\beta\beta$  decay of  $^{160}\text{Gd}$  into excited  $2^+$  state of daughter nucleus using  $\text{Gd}_2\text{SiO}_5\text{:Ce}$  crystal scintillators.

<sup>73</sup> KLAPDOR-KLEINGROTHAUS 01 is a continuation of the work published in BAUDIS 99. Isotopically enriched Ge detectors are used in calorimetric measurement. The most stringent bound is derived from the data set in which pulse-shape analysis has been used to reduce background. Exposure time is 35.5 kg y. Supersedes BAUDIS 99 as most stringent result.

<sup>74</sup> WIESER 01 reports an inclusive geochemical measurement of  $^{96}\text{Zr}$   $\beta\beta$  half life. Their result agrees within  $2\sigma$  with ARNOLD 99 but only marginally, within  $3\sigma$ , with KAWASHIMA 93.

<sup>75</sup> BRUDANIN 00 determine the  $2\nu\beta\beta$  half-life of  $^{48}\text{Ca}$ . Their value is less accurate than BALYSH 96.

<sup>76</sup> ARNOLD 99 measure directly the  $2\nu\beta\beta$  decay of Zr for the first time, using the NEMO-2 tracking detector and an isotopically enriched source. The lifetime is more accurate than the geochemical result of KAWASHIMA 93.

<sup>77</sup> ARNOLD 98 determine the limit for  $0\nu\beta\beta$  decay to the excited  $2^+$  state of  $^{82}\text{Se}$  using the NEMO-2 tracking detector.

<sup>78</sup> DESILVA 97 result for  $2\nu\beta\beta$  decay of  $^{150}\text{Nd}$  is in marginal agreement with ARTEMEV 93. It has smaller errors.

<sup>79</sup> BALYSH 96 measure the  $2\nu\beta\beta$  decay of  $^{48}\text{Ca}$ , using a passive source of enriched  $^{48}\text{Ca}$  in a TPC.

<sup>80</sup> BERNATOWICZ 92 finds  $^{128}\text{Te}/^{130}\text{Te}$  activity ratio from slope of  $^{128}\text{Xe}/^{132}\text{Xe}$  vs  $^{130}\text{Xe}/^{132}\text{Xe}$  ratios during extraction, and normalizes to lead-dated ages for the  $^{130}\text{Te}$  lifetime. The authors state that their results imply that “(a) the double beta decay of  $^{128}\text{Te}$  has been firmly established and its half-life has been determined ... without any ambiguity due to trapped Xe interferences. ... (b) Theoretical calculations ... underestimate the [long half-lives of  $^{128}\text{Te}$   $^{130}\text{Te}$ ] by 1 or 2 orders of magnitude, pointing to a real suppression in the  $2\nu\beta\beta$  decay rate of these isotopes. (c) Despite [this], most  $\beta\beta$ -models predict a ratio of  $2\nu\beta\beta$  decay widths ... in fair agreement with observation.” Further details of the experiment are given in BERNATOWICZ 93. Our listed half-life has been revised downward from the published value by the authors, on the basis of reevaluated cosmic-ray  $^{128}\text{Xe}$  production corrections.

<sup>81</sup> TURKEVICH 91 observes activity in old U sample. The authors compare their results with theoretical calculations. They state “Using the phase-space factors of Boehm and Vogel (BOEHM 87) leads to matrix element values for the  $^{238}\text{U}$  transition in the same range as deduced for  $^{130}\text{Te}$  and  $^{76}\text{Ge}$ . On the other hand, the latest theoretical estimates (STAUDT 90) give an upper limit that is 10 times lower. This large discrepancy implies either a defect in the calculations or the presence of a faster path than the standard two-neutrino mode in this case.” See BOEHM 87 and STAUDT 90.

<sup>82</sup> Ratio of inclusive double beta half lives of  $^{128}\text{Te}$  and  $^{130}\text{Te}$  determined from minerals melonite ( $\text{NiTe}_2$ ) and altaite ( $\text{PbTe}$ ) by means of mass spectroscopic measurement of abundance of  $\beta\beta$ -decay products. As gas-retention-age could not be determined the authors use half life of  $^{130}\text{Te}$  (LIN 88) to infer the half life of  $^{128}\text{Te}$ . No estimate of the systematic uncertainty of this method is given. The directly determined half life ratio agrees with BERNATOWICZ 92. However, the inferred  $^{128}\text{Te}$  half life disagrees with KIRSTEN 83 and BERNATOWICZ 92.

### $\langle m_\nu \rangle$ , The Effective Weighted Sum of Majorana Neutrino Masses Contributing to Neutrinoless Double-β Decay

$\langle m_\nu \rangle = |\sum U_{ej}^2 m_\nu|$ , where the sum goes from 1 to  $n$  and where  $n$  = number of neutrino generations, and  $\nu_j$  is a Majorana neutrino. Note that  $U_{ej}^2$ , not  $|U_{ej}|^2$ , occurs in the sum. The possibility of cancellations has been stressed. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported.

VALUE (eV)	CL% ISOTOPE	TRANSITION	METHOD	DOCUMENT ID
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
< 0.27–0.65	90 $^{130}\text{Te}$	$0\nu, \text{g.s.} \rightarrow \text{g.s.}$	CUORE	1 ALFONSO 15
< 0.33–0.62	90 $^{100}\text{Mo}$	$0\nu$	NEMO-3	2 ARNOLD 15
< 0.19–0.45	90 $^{136}\text{Xe}$	$0\nu, \text{g.s.} \rightarrow \text{g.s.}$	EXO-200	3 ALBERT 14b
< 0.2–0.4	90 $^{76}\text{Ge}$	$0\nu$	GERDA	4 AGOSTINI 13A
< 0.12–0.25	90 $^{136}\text{Xe}$	$0\nu, \text{g.s.} \rightarrow \text{g.s.}$	KamLAND-Zen	5 GANDO 13A
< 0.3–0.6	90 $^{136}\text{Xe}$	$0\nu, \text{g.s.} \rightarrow \text{g.s.}$	KamLAND-Zen	6 GANDO 12A
< 0.89–2.43	90 $^{82}\text{Se}$	$0\nu$	NEMO-3	7 BARABASH 11A
< 7.2–19.5	90 $^{96}\text{Zr}$	$0\nu$	NEMO-3	8 ARGYRIADES 10
< 4.0–6.8	90 $^{150}\text{Nd}$	$0\nu$	NEMO-3	9 ARGYRIADES 09

# Lepton Particle Listings

## Double- $\beta$ Decay

< 3.5–22	90	<sup>48</sup> Ca	0 $\nu$	CaF <sub>2</sub> scint.	10	UMEHARA	08
< 9.3–60	90	<sup>100</sup> Mo	0 $^+ \rightarrow 0^+_{1/2}$	NEMO-3	11	ARNOLD	07
< 6500	90	<sup>100</sup> Mo	0 $^+ \rightarrow 2^+$	NEMO-3	12	ARNOLD	07
0.32 $\pm$ 0.03	68	<sup>76</sup> Ge	0 $\nu$	Enriched HPGe	13	KLAPDOR-K...	06A
< 0.2–1.1	90	<sup>130</sup> Te		Cryog. det.	14	ARNABOLDI	05
< 0.7–2.8	90	<sup>100</sup> Mo	0 $\nu$	NEMO-3	15	ARNOLD	05A
< 1.7–4.9	90	<sup>82</sup> Se	0 $\nu$	NEMO-3	16	ARNOLD	05A
< 0.37–1.9	90	<sup>130</sup> Te		Cryog. det.	17	ARNABOLDI	04
< 0.8–1.2	90	<sup>100</sup> Mo	0 $\nu$	NEMO-3	18	ARNOLD	04
< 1.5–3.1	90	<sup>82</sup> Se	0 $\nu$	NEMO-3	18	ARNOLD	04
0.1–0.9	99.7	<sup>76</sup> Ge		Enriched HP Ge	19	KLAPDOR-K...	04A
< 7.2–44.7	90	<sup>48</sup> Ca		CaF <sub>2</sub> scint.	20	OGAWA	04
< 1.1–2.6	90	<sup>130</sup> Te		Cryog. det.	21	ARNABOLDI	03
< 1.5–1.7	90	<sup>116</sup> Cd	0 $\nu$	<sup>116</sup> CdWO <sub>4</sub> scint.	22	DANEVICH	03
< 0.33–1.35	90			Enriched HPGe	23	AALSETH	02B
< 2.9	90	<sup>136</sup> Xe	0 $\nu$	Liquid Xe Scint.	24	BERNABEI	02D
0.39 $\pm$ 0.17 –0.28		<sup>76</sup> Ge	0 $\nu$	Enriched HPGe	25	KLAPDOR-K...	02D
< 2.1–4.8	90	<sup>100</sup> Mo	0 $\nu$	ELEGANT V	26	EJIRI	01
< 0.35	90	<sup>76</sup> Ge		Enriched HPGe	27	KLAPDOR-K...	01
< 23	90	<sup>96</sup> Zr		NEMO-2	28	ARNOLD	99
< 1.1–1.5		<sup>128</sup> Te		Geochem	29	BERNATOW...	92
< 5	68	<sup>82</sup> Se		TPC	30	ELLIOTT	92
< 8.3	76	<sup>48</sup> Ca	0 $\nu$	CaF <sub>2</sub> scint.		YOU	91

- ALFONSO 15 report a range of mass limits using the combined data of the CUORICINO and CUORE-0 experiments. The reported mass range reflects the variability of the nuclear matrix element calculations.
- ARNOLD 15 use the NEMO-3 tracking calorimeter with 34.3 kg yr exposure to determine the neutrino mass limit based on the 0 $\nu\beta\beta$ -half life of <sup>100</sup>Mo. The spread range reflects different nuclear matrix elements. Supersedes ARNOLD 14 and BARABASH 11A.
- ALBERT 14B is based on 100 kg yr of exposure of the EXO-200 tracking calorimeter. The mass range reflects the nuclear matrix element calculations. Supersedes AUGER 12.
- AGOSTINI 13A is based on 21.6 kg yr of data collected by the GERDA detector. The reported range reflects different nuclear matrix elements. This result is in tension with the evidence for 0 $\nu\beta\beta$ -decay reported in KLAPDOR-KLEINGROTHAUS 06A and earlier references to that work.
- GANDO 13A limit is based on a combination of KamLAND-Zen and EXO-200 (AUGER 12) data. The reported range reflects different nuclear matrix elements. Supersedes GANDO 12A.
- GANDO 12A limit is based on the KamLAND-Zen data. The reported range reflects different nuclear matrix elements. Superseded by GANDO 13A.
- BARABASH 11A limit is based on NEMO-3 data for <sup>82</sup>Se. The reported range reflects different nuclear matrix elements. Supersedes ARNOLD 05A and ARNOLD 04.
- ARGYRIADES 10 use <sup>96</sup>Zr and the NEMO-3 tracking detector to obtain the reported mass limit. The range reflects the fluctuation of the nuclear matrix elements considered.
- ARGYRIADES 09 limit is based on data taken with the NEMO-3 detector and <sup>150</sup>Nd. A range of nuclear matrix elements that include the effect of nuclear deformation have been used.
- Limit was obtained using CaF<sub>2</sub> scintillation calorimeter to search for double beta decay of <sup>48</sup>Ca. Reported range of limits reflects spread of QRPA and SM matrix element calculations used. Supersedes OGAWA 04.
- ARNOLD 07 use NEMO-3 half life limit for 0 $\nu$ -decay of <sup>100</sup>Mo to the first excited 0 $^+_{1/2}$ -state of daughter nucleus to obtain neutrino mass limit. The spread reflects the choice of two different nuclear matrix elements. This limit is not competitive when compared to the decay to the ground state.
- ARNOLD 07 use NEMO-3 half life limit for 0 $\nu$ -decay of <sup>100</sup>Mo to the first excited 2 $^+_{1/2}$ -state of daughter nucleus to obtain neutrino mass limit. This limit is not competitive when compared to the decay to the ground state.
- Re-analysis of data originally published in KLAPDOR-KLEINGROTHAUS 04A. Modified pulse shape analysis leads the authors to claim 6 $\sigma$  statistical evidence for observation of 0 $\nu$ -decay. Authors use matrix element of STAUDT 90. Uncertainty of nuclear matrix element is not reflected in stated error. Supersedes KLAPDOR-KLEINGROTHAUS 04A.
- Supersedes ARNABOLDI 04. Reported range of limits due to use of different nuclear matrix element calculations.
- Mass limits reported in ARNOLD 05A are derived from <sup>100</sup>Mo data, obtained by the NEMO-3 collaboration. The range reflects the spread of matrix element calculations considered in this work. Supersedes ARNOLD 04.
- Neutrino mass limits based on <sup>82</sup>Se data utilizing the NEMO-3 detector. The range reported in ARNOLD 05A reflects the spread of matrix element calculations considered in this work. Supersedes ARNOLD 04.
- Supersedes ARNABOLDI 03. Reported range of limits due to use of different nuclear matrix element calculations.
- ARNOLD 04 limit is based on the nuclear matrix elements of SIMKOVIC 99, STOICA 01 and CIVITARESE 03.
- Supersedes KLAPDOR-KLEINGROTHAUS 02D. Event excess at  $\beta\beta$ -decay energy is used to derive Majorana neutrino mass using the nuclear matrix elements of STAUDT 90. The mass range shown is based on the authors evaluation of the uncertainties of the STAUDT 90 matrix element calculation. If this uncertainty is neglected, and only statistical errors are considered, the range in  $\langle m \rangle$  becomes (0.2–0.6) eV at the 3  $\sigma$  level.
- Calorimetric CaF<sub>2</sub> scintillator. Range of limits reflects authors' estimate of the uncertainty of the nuclear matrix elements. Replaces YOU 91 as the most stringent limit based on <sup>48</sup>Ca.
- Supersedes ALESSANDRELLO 00. Cryogenic calorimeter search. Reported a range reflecting uncertainty in nuclear matrix element calculations.

- Limit for  $\langle m_{\nu} \rangle$  is based on the nuclear matrix elements of STAUDT 90 and ARNOLD 96. Supersedes DANEVICH 00.
- AALSETH 02B reported range of limits on  $\langle m_{\nu} \rangle$  reflects the spread of theoretical nuclear matrix elements. Excludes part of allowed mass range reported in KLAPDOR-KLEINGROTHAUS 01B.
- BERNABEI 02D limit is based on the matrix elements of SIMKOVIC 02. The range of neutrino masses based on a variety of matrix elements is 1.1–2.9 eV.
- KLAPDOR-KLEINGROTHAUS 02D is a detailed description of the analysis of the data collected by the Heidelberg-Moscow experiment, previously presented in KLAPDOR-KLEINGROTHAUS 01B. Matrix elements in STAUDT 90 have been used. See the footnote in the preceding table for further details. See also KLAPDOR-KLEINGROTHAUS 02B.
- The range of the reported  $\langle m_{\nu} \rangle$  values reflects the spread of the nuclear matrix elements. On axis value assuming  $\langle \lambda \rangle = \langle \eta \rangle = 0$ .
- KLAPDOR-KLEINGROTHAUS 01 uses the calculation by STAUDT 90. Using several other models in the literature could worsen the limit up to 1.2 eV. This is the most stringent experimental bound on  $m_{\nu}$ . It supersedes BAUDIS 99B.
- ARNOLD 99 limit based on the nuclear matrix elements of STAUDT 90.
- BERNATOWICZ 92 finds these majorana neutrino mass limits assuming that the measured geochemical decay width is a limit on the 0 $\nu$  decay width. The range is the range found using matrix elements from HAXTON 84, TOMODA 87, and SUHONEN 91. Further details of the experiment are given in BERNATOWICZ 93.
- ELLIOTT 92 uses the matrix elements of HAXTON 84.

### Limits on Lepton-Number Violating (V+A) Current Admixture

For reasons given in the discussion at the beginning of this section, we list only results from 1989 and later.  $\langle \lambda \rangle = \lambda \sum U_{ej} V_{ej}$  and  $\langle \eta \rangle = \eta \sum U_{ej} V_{ej}$ , where the sum is over the number of neutrino generations. This sum vanishes for massless or unmixed neutrinos. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported.

$\langle \lambda \rangle$ ( $10^{-6}$ )	CL%	$\langle \eta \rangle$ ( $10^{-8}$ )	CL%	ISOTOPE	METHOD	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •						
< 0.9–1.3	90	< 0.5–0.8	90	<sup>100</sup> Mo	NEMO-3	1 ARNOLD 14
< 120	90			<sup>100</sup> Mo	0 $^+ \rightarrow 2^+$	2 ARNOLD 07
0.692 $\pm$ 0.058 –0.056	68	0.305 $\pm$ 0.026 –0.025	68	<sup>76</sup> Ge	Enriched HPGe	3 KLAPDOR-K... 06A
< 2.5	90			<sup>100</sup> Mo	0 $\nu$ , NEMO-3	4 ARNOLD 05A
< 3.8	90			<sup>82</sup> Se	0 $\nu$ , NEMO-3	5 ARNOLD 05A
< 1.5–2.0	90			<sup>100</sup> Mo	0 $\nu$ , NEMO-3	6 ARNOLD 04
< 3.2–3.8	90			<sup>82</sup> Se	0 $\nu$ , NEMO-3	7 ARNOLD 04
< 1.6–2.4	90	< 0.9–5.3	90	<sup>130</sup> Te	Cryog. det.	8 ARNABOLDI 03
< 2.2	90	< 2.5	90	<sup>116</sup> Cd	<sup>116</sup> CdWO <sub>4</sub> scint.	9 DANEVICH 03
< 3.2–4.7	90	< 2.4–2.7	90	<sup>100</sup> Mo	ELEGANT V	10 EJIRI 01
< 1.1	90	< 0.64	90	<sup>76</sup> Ge	Enriched HPGe	11 GUENTHER 97
< 4.4	90	< 2.3	90	<sup>136</sup> Xe	TPC	12 VUILLEUMIER 93
		< 5.3		<sup>128</sup> Te	Geochem	13 BERNATOW... 92

- ARNOLD 14 is based on 34.7 kg yr of exposure of the NEMO-3 tracking calorimeter. The reported range limit on  $\langle \lambda \rangle$  and  $\langle \eta \rangle$  reflects the nuclear matrix element uncertainty in <sup>100</sup>Mo.
- ARNOLD 07 use NEMO-3 half life limit for 0 $\nu$ -decay of <sup>100</sup>Mo to the first excited 2 $^+_{1/2}$ -state of daughter nucleus to limit the right-right handed admixture of weak currents  $\langle \lambda \rangle$ . This limit is not competitive when compared to the decay to the ground state.
- Re-analysis of data originally published in KLAPDOR-KLEINGROTHAUS 04A. Modified pulse shape analysis leads the authors to claim 6 $\sigma$  statistical evidence for observation of 0 $\nu$ -decay. Authors use matrix element of MUTO 89 to determine  $\langle \lambda \rangle$  and  $\langle \eta \rangle$ . Uncertainty of nuclear matrix element is not reflected in stated errors.
- ARNOLD 05A derive limit for  $\langle \lambda \rangle$  based on <sup>100</sup>Mo data collected with NEMO-3 detector. No limit for  $\langle \eta \rangle$  is given. Supersedes ARNOLD 04.
- ARNOLD 05A derive limit for  $\langle \lambda \rangle$  based on <sup>82</sup>Se data collected with NEMO-3 detector. No limit for  $\langle \eta \rangle$  is given. Supersedes ARNOLD 04.
- ARNOLD 04 use the matrix elements of SUHONEN 94 to obtain a limit for  $\langle \lambda \rangle$ , no limit for  $\langle \eta \rangle$  is given. This limit is more stringent than the limit in EJIRI 01 for the same nucleus.
- ARNOLD 04 use the matrix elements of TOMODA 91 and SUHONEN 91 to obtain a limit for  $\langle \lambda \rangle$ , no limit for  $\langle \eta \rangle$  is given.
- Supersedes ALESSANDRELLO 00. Cryogenic calorimeter search. Reported a range reflecting uncertainty in nuclear matrix element calculations.
- Limits for  $\langle \lambda \rangle$  and  $\langle \eta \rangle$  are based on nuclear matrix elements of STAUDT 90. Supersedes DANEVICH 00.
- The range of the reported  $\langle \lambda \rangle$  and  $\langle \eta \rangle$  values reflects the spread of the nuclear matrix elements. On axis value assuming  $\langle m_{\nu} \rangle = 0$  and  $\langle \lambda \rangle = \langle \eta \rangle = 0$ , respectively.
- GUENTHER 97 limits use the matrix elements of STAUDT 90. Supersedes BALYSH 95 and BALYSH 92.
- VUILLEUMIER 93 uses the matrix elements of MUTO 89. Based on a half-life limit  $2.6 \times 10^{23}$  y at 90%CL.
- BERNATOWICZ 92 takes the measured geochemical decay width as a limit on the 0 $\nu$  width, and uses the SUHONEN 91 coefficients to obtain the least restrictive limit on  $\eta$ . Further details of the experiment are given in BERNATOWICZ 93.

# Lepton Particle Listings

## Double- $\beta$ Decay, Neutrino Mixing

### Double- $\beta$ Decay REFERENCES

ASAKURA	16	NP A946 171	K. Asakura <i>et al.</i>	(KamLAND-Zen Collab.)
AGOSTINI	15A	EPJ C75 416	M. Agostini <i>et al.</i>	(GERDA Collab.)
ALFONSO	15	PRL 115 102502	K. Alfonso <i>et al.</i>	(CUORE Collab.)
ARNOLD	15	PR D92 072011	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
ALBERT	14	PR C89 015502	J. Albert <i>et al.</i>	(EXO-200 Collab.)
ALBERT	14B	NAT 510 229	J.B. Albert <i>et al.</i>	(EXO-200 Collab.)
ARNOLD	14	PR D89 111101	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
KIDD	14	PR C90 055501	M.F. Kidd <i>et al.</i>	
AGOSTINI	13A	PRL 111 122503	M. Agostini <i>et al.</i>	(GERDA Collab.)
BELLI	13A	PR C87 034607	P. Belli <i>et al.</i>	(DAMA-INR Collab.)
GANDO	13A	PRL 110 062502	A. Gando <i>et al.</i>	(KamLAND-Zen Collab.)
GAVRILYAK	13	PR C87 035501	Yu.M. Gavriluk <i>et al.</i>	
SCHWINGEN...	13	ANP 525 269	B. Schwingenheuer	(MPIH)
ANDREOTTI	12	PR C85 045503	E. Andreotti <i>et al.</i>	(CUORICINO Collab.)
AUGER	12	PRL 109 032505	M. Auger <i>et al.</i>	(EXO-200 Collab.)
BELLI	12A	PR C85 044610	P. Belli <i>et al.</i>	
GANDO	12A	PR C85 045504	A. Gando <i>et al.</i>	(KamLAND-Zen Collab.)
ACKERMAN	11	PRL 107 212501	N. Ackerman <i>et al.</i>	(EXO Collab.)
ARNOLD	11	PRL 107 062504	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
BARABASH	11	PR C83 045503	A.S. Barabash <i>et al.</i>	
BARABASH	11A	PAN 74 312	A.S. Barabash <i>et al.</i>	(NEMO-3 Collab.)
BELLI	11D	Translated from YAF 74 330		
RUKHADZE	11	JP G38 115107	P. Belli <i>et al.</i>	(DAMA-INR Collab.)
ARGYRIADES	10	NP A852 197	A. Rukhadze <i>et al.</i>	(TGV-2 Collab.)
BELLI	10	NP A847 168	J. Argyriades <i>et al.</i>	(NEMO-3 Collab.)
BELLI	10	NP A845 143	P. Belli <i>et al.</i>	(DAMA-INR Collab.)
ARGYRIADES	09	PR C80 032501	J. Argyriades <i>et al.</i>	(NEMO-3 Collab.)
BELLI	09A	NP A826 256	P. Belli <i>et al.</i>	(DAMA-INR Collab.)
KIDD	09	NP A821 251	M. Kidd <i>et al.</i>	
ARNABOLDI	08	PR C78 035502	C. Arnaboldi <i>et al.</i>	
BELLI	08	PL B658 193	P. Belli <i>et al.</i>	(DAMA-INR Collab.)
BELLI	08B	EPJ A36 167	P. Belli <i>et al.</i>	
UMEHARA	08	PR C78 058501	S. Umehara <i>et al.</i>	
ARNOLD	07	NP A781 209	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
KLAPDOR-K...	06A	MPL A21 1547	H.V. Klapdor-Kleingrothaus, I.V. Krivosheina	
ARNABOLDI	05	PRL 95 142501	C. Arnaboldi <i>et al.</i>	(CUORICINO Collab.)
ARNOLD	05A	PRL 95 182302	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
AALSETH	04	PR D70 078302	C.E. Aalseth <i>et al.</i>	
ARNABOLDI	04	PL B584 260	C. Arnaboldi <i>et al.</i>	
ARNOLD	04	JETPL 80 377	R. Arnold <i>et al.</i>	(NEMO-3 Collab.)
BARABASH	04	Translated from ZETFP 80 429		
KLAPDOR-K...	04A	JETPL 79 10	A.S. Barabash <i>et al.</i>	
KLAPDOR-K...	04B	PL B586 198	H.V. Klapdor-Kleingrothaus <i>et al.</i>	
OGAWA	04	PR D70 078301	H.V. Klapdor-Kleingrothaus, A. Dietz, I.V. Krivosheina	
ARNABOLDI	03	NP A730 215	I. Ogawa <i>et al.</i>	
CIVITARESE	03	PL B557 167	C. Arnaboldi <i>et al.</i>	
DANEVICH	03	NP A729 867	O. Civitarese, J. Suhonen	
AALSETH	03	PR C68 035501	F.A. Danevich <i>et al.</i>	(IGEX Collab.)
BERNABEI	02D	PR D65 092007	C.E. Aalseth <i>et al.</i>	
KLAPDOR-K...	02B	PL B546 23	R. Bernabei <i>et al.</i>	(DAMA Collab.)
KLAPDOR-K...	02D	PPNL 110 57	H.V. Klapdor-Kleingrothaus, A. Dietz, I.V. Krivosheina	
KLAPDOR-K...	02D	FP 32 1181	H.V. Klapdor-Kleingrothaus, A. Dietz, I.V. Krivosheina	
SIMKOVIC	02	hep-ph/0204278	F. Simkovic, P. Domin, A. Faessler	
DANEVICH	01	NP A694 375	F.A. Danevich <i>et al.</i>	
DEBRAECKEL...	01	PRL 86 3510	L. De Braeckeleer <i>et al.</i>	
EJIRI	01	PR C63 065501	H. Ejiri <i>et al.</i>	
KLAPDOR-K...	01	EPJ A12 147	H.V. Klapdor-Kleingrothaus <i>et al.</i>	
KLAPDOR-K...	01B	MPL A16 2409	H.V. Klapdor-Kleingrothaus <i>et al.</i>	
STOICA	01	NP A694 269	S. Stoica, H.V. Klapdor-Kleingrothaus	
WIESSER	01	PR C64 024308	M.E. Wieser, J.R. De Laeter	
ALESSANDRO...	00	PL B486 13	A. Alessandrello <i>et al.</i>	
BRUDANIN	00	PL B495 63	V.B. Brudanin <i>et al.</i>	(TGV Collab.)
DANEVICH	00	PR C62 045501	F.A. Danevich <i>et al.</i>	
ARNOLD	99	NP A658 299	R. Arnold <i>et al.</i>	(NEMO Collab.)
BAUDIS	99	PR D59 022001	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
BAUDIS	99B	PRL 83 41	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
SIMKOVIC	99	PR C60 055502	F. Simkovic <i>et al.</i>	
ARNOLD	98	NP A636 209	R. Arnold <i>et al.</i>	(NEMO-2 Collab.)
DESILVA	97	PR C56 2451	A. de Silva <i>et al.</i>	(UCI)
GUNTHER	97	PR D55 54	M. Gunther <i>et al.</i>	(Heidelberg-Moscow Collab.)
ARNOLD	96	ZPHY C72 239	R. Arnold <i>et al.</i>	(BCEH, CAEN, JINR+)
BALYSH	96	PRL 77 5186	A. Balysh <i>et al.</i>	(KIAE, UCI, CIT)
ARNOLD	95	JETPL 61 170	R.G. Arnold <i>et al.</i>	(NEMO Collab.)
BALYSH	95	Translated from ZETFP 61 168		
BARABASH	95	PL B356 450	A. Balysh <i>et al.</i>	(Heidelberg-Moscow Collab.)
DASSIE	95	PL B345 408	A.S. Barabash <i>et al.</i>	(ITEP, SCUC, PNL+)
EJIRI	95	PR D51 2090	D. Dassie <i>et al.</i>	(NEMO Collab.)
KOBAYASHI	95	JPJ 4 64 339	H. Ejiri <i>et al.</i>	(OSAK, KIEV)
SUHONEN	94	NP A586 457	M. Kobayashi, M. Kobayashi	(KEK, SAGA)
ARTEMEYEV	93	PR C49 3055	J. Suhonen, O. Civitarese	
BERNATOW...	93	JETPL 58 262	V.A. Artemiev <i>et al.</i>	(ITEP, INRM)
KAWASHIMA	93	Translated from ZETFP 58 256		
VUILLEUMIER	93	PR C47 806	T. Bernatowicz <i>et al.</i>	(WUSL, TATA)
BALYSH	93	PR C47 R2452	A. Kawashima, K. Takahashi, A. Masuda	(TOKYU+)
ELLIOTT	92	PL D48 1009	J.C. Vuilleumier <i>et al.</i>	(NEUC, CIT, VILL)
BERNATOW...	92	PL B283 32	A. Balysh <i>et al.</i>	(MPIH, KIAE, SASSO)
SUHONEN	91	PRL 69 2341	T. Bernatowicz <i>et al.</i>	(WUSL, TATA)
TOMODA	91	PR C46 1535	S.R. Elliott <i>et al.</i>	(UCI)
TURKEVICH	91	NP A535 509	J. Suhonen, S.B. Khadikar, A. Faessler	(JYV+)
YOU	91	RPP 54 53	T. Tomoda	
STAUDT	90	PRL 67 3211	A. Turkevich, T.E. Economou, G.A. Cowan	(CHIC+)
MUTO	89	PL B265 53	K. You <i>et al.</i>	(BHEP, CAST+)
LIN	88	EPL 13 31	A. Staudt, K. Muto, H.V. Klapdor-Kleingrothaus	(TINT, MPIH)
BOEHM	88B	ZPHY A334 187	K. Muto, E. Bender, H.V. Klapdor	
BOEHM	87	NP A481 477	W.J. Lin <i>et al.</i>	
BOEHM	87	NP A481 484	W.J. Lin <i>et al.</i>	
BOEHM	87	Massive Neutrinos	F. Bohm, P. Vogel	(CIT)
TOMODA	87	Cambridge Univ. Press, Cambridge		
HAXTON	87	PL B199 475	T. Tomoda, A. Faessler	(TUBIN)
KIRSTEN	83	PPNP 12 409	W.C. Haxton, G.J. Stevenson	
KIRSTEN	83	PRL 50 474	T. Kirsten, H. Richter, E. Jessberger	(MPIH)

## Neutrino Mixing

With the exception of a few possible anomalies such as LSND, current neutrino data can be described within the framework of a  $3\times 3$  mixing matrix between the flavor eigenstates  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$  and the mass eigenstates  $\nu_1$ ,  $\nu_2$ , and

$\nu_3$ . (See Eq. (14.6) of the review “Neutrino Mass, Mixing, and Oscillations” by K. Nakamura and S.T. Petcov.) The Listings are divided into the following sections:

**(A) Neutrino fluxes and event ratios:** shows measurements which correspond to various oscillation tests for Accelerator, Reactor, Atmospheric, and Solar neutrino experiments. Typically ratios involve a measurement in a realm sensitive to oscillations compared to one for which no oscillation effect is expected.

**(B) Three neutrino mixing parameters:** shows measurements of  $\sin^2(2\theta_{12})$ ,  $\sin^2(2\theta_{23})$ ,  $\Delta m_{21}^2$ ,  $\Delta m_{32}^2$ , and  $\sin^2(2\theta_{13})$  which are all interpretations of data based on the three neutrino mixing scheme described in the review “Neutrino Mass, Mixing, and Oscillations.” by K. Nakamura and S.T. Petcov. Many parameters have been calculated in the two-neutrino approximation.

**(C) Other neutrino mixing results:** shows measurements and limits for the probability of oscillation for experiments which might be relevant to the LSND oscillation claim. Included are experiments which are sensitive to  $\nu_\mu \rightarrow \nu_e$ ,  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ , sterile neutrinos, and CPT tests.

### (A) Neutrino fluxes and event ratios

#### Events (observed/expected) from accelerator $\nu_\mu$ experiments.

Some neutrino oscillation experiments compare the flux in two or more detectors. This is usually quoted as the ratio of the event rate in the far detector to the expected rate based on an extrapolation from the near detector in the absence of oscillations.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.01 ± 0.10	<sup>1</sup> ABE	14B T2K	$\nu_e$ rate in T2K near detect.
0.71 ± 0.08	<sup>2</sup> AHN	06A K2K	K2K to Super-K
0.64 ± 0.05	<sup>3</sup> MICHAEL	06 MINS	All charged current events
0.71 +0.08 -0.09	<sup>4</sup> ALIU	05 K2K	KEK to Super-K
0.70 +0.10 -0.11	<sup>5</sup> AHN	03 K2K	KEK to Super-K

<sup>1</sup> The rate of  $\nu_e$  from  $\mu$  decay was measured to be  $0.68 \pm 0.30$  compared to the predicted flux. From  $K$  decay  $1.10 \pm 0.14$  compared to the predicted flux.

<sup>2</sup> Based on the observation of 112 events when  $158.1^{+9.2}_{-8.6}$  were expected without oscillations. Including not only the number of events but also the shape of the energy distribution, the evidence for oscillation is at the level of about  $4.3 \sigma$ . Supersedes ALIU 05.

<sup>3</sup> This ratio is based on the observation of 215 events compared to an expectation of  $336 \pm 14$  without oscillations. See also ADAMSON 08.

<sup>4</sup> This ratio is based on the observation of 107 events at the far detector 250 km away from KEK, and an expectation of  $151^{+12}_{-10}$ .

<sup>5</sup> This ratio is based on the observation of 56 events with an expectation of  $80.1^{+6.2}_{-5.4}$ .

#### Events (observed/expected) from reactor $\bar{\nu}_e$ experiments.

The quoted values are the ratios of the measured reactor  $\bar{\nu}_e$  event rate at the quoted distances, and the rate expected without oscillations. The expected rate is based on the experimental data for the most significant reactor fuels ( $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$ ) and on calculations for  $^{238}\text{U}$ .

A recent re-evaluation of the spectral conversion of electron to  $\bar{\nu}_e$  in MUELLER 11 results in an upward shift of the reactor  $\bar{\nu}_e$  spectrum by 3% and, thus, might require revisions to the ratios listed in this table.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.944 ± 0.007 ± 0.003</b>	<sup>1</sup> AN	13 DAYA	DayaBay, Ling Ao/Ao II reactors
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.944 ± 0.016 ± 0.040	<sup>2</sup> ABE	12 DCHZ	Chooz reactors
0.920 ± 0.009 ± 0.014	<sup>3</sup> AHN	12 RENO	Yonggwang reactors
0.940 ± 0.011 ± 0.004	<sup>4</sup> AN	12 DAYA	DayaBay, Ling Ao/Ao II reactors
1.08 ± 0.21 ± 0.16	<sup>5</sup> DENIZ	10 TEXO	Kuo-Sheng reactor, 28 m
0.658 ± 0.044 ± 0.047	<sup>6</sup> ARAKI	05 KLND	Japanese react. ~180 km
0.611 ± 0.085 ± 0.041	<sup>7</sup> EGUCHI	03 KLND	Japanese react. ~180 km
1.01 ± 0.024 ± 0.053	<sup>8</sup> BOEHM	01	Palo Verde react. 0.75–0.89 km
1.01 ± 0.028 ± 0.027	<sup>9</sup> APOLLONIO	99 CHOZ	Chooz reactors 1 km

0.987±0.006±0.037	<sup>10</sup> GREENWOOD 96	Savannah River, 18.2 m
0.988±0.004±0.05	ACHKAR 95	CNTR Bugey reactor, 15 m
0.994±0.010±0.05	ACHKAR 95	CNTR Bugey reactor, 40 m
0.915±0.132±0.05	ACHKAR 95	CNTR Bugey reactor, 95 m
0.987±0.014±0.027	<sup>11</sup> DECLAIS 94	CNTR Bugey reactor, 15 m
0.985±0.018±0.034	KUVSHINN... 91	CNTR Rovno reactor
1.05 ±0.02 ±0.05	VUILLEUMIER 82	Gösgen reactor
0.955±0.035±0.110	<sup>12</sup> KWON 81	$\bar{\nu}_e p \rightarrow e^+ n$
0.89 ±0.15	<sup>12</sup> BOEHM 80	$\bar{\nu}_e p \rightarrow e^+ n$

<sup>1</sup> AN 13 use six identical detectors, with three placed near the reactor cores (flux-weighted baselines of 470 and 576 m) and the remaining three at the far hall (at the flux averaged distance of 1648 m from all six reactor cores) to determine the mixing angle  $\theta_{13}$  using the  $\bar{\nu}_e$  observed interaction rate ratios. This rate-only analysis excludes the no-oscillation hypothesis at 7.7 standard deviations. The value of  $\Delta m_{31}^2 = 2.32 \times 10^{-3} \text{ eV}^2$  was assumed in the analysis. This is an improved result (2.5 times increase in statistics) compared to AN 12.

<sup>2</sup> ABE 12 determine the  $\bar{\nu}_e$  interaction rate in a single detector, located 1050 m from the cores of two reactors. The rate normalization is fixed by the results of the Bugey4 reactor experiment, thus avoiding any dependence on possible very short baseline oscillations.

<sup>3</sup> AHN 12 use two identical detectors, placed at flux weighted distances of 408.56 m and 1433.99m from six reactor cores, to determine the  $\bar{\nu}_e$  interaction rate ratio.

<sup>4</sup> AN 12 use six identical detectors with three placed near the reactor cores (flux-weighted baselines of 470 m and 576 m) and the remaining three at the far hall (at the flux averaged distance of 1648 m from all six reactor cores) to determine the  $\bar{\nu}_e$  interaction rate ratios. Superseded by AN 13.

<sup>5</sup> DENIZ 10 observe reactor  $\bar{\nu}_e e$  scattering with recoil kinetic energies 3–8 MeV using CsI(Tl) detectors. The observed rate is consistent with the Standard Model prediction, leading to a constraint on  $\sin^2 2\theta_{\text{eff}} = 0.251 \pm 0.031(\text{stat}) \pm 0.024(\text{sys})$ .

<sup>6</sup> Updated result of KamLAND, including the data used in EGUCHI 03. Note that the survival probabilities for different periods are not directly comparable because the effective baseline varies with power output of the reactor sources involved, and there were large variations in the reactor power production in Japan in 2003.

<sup>7</sup> EGUCHI 03 observe reactor neutrino disappearance at  $\sim 180 \text{ km}$  baseline to various Japanese nuclear power reactors.

<sup>8</sup> BOEHM 01 search for neutrino oscillations at 0.75 and 0.89 km distance from the Palo Verde reactors.

<sup>9</sup> APOLLONIO 99, APOLLONIO 98 search for neutrino oscillations at 1.1 km fixed distance from Chooz reactors. They use  $\bar{\nu}_e p \rightarrow e^+ n$  in Gd-loaded scintillator target. APOLLONIO 99 supersedes APOLLONIO 98. See also APOLLONIO 03 for detailed description.

<sup>10</sup> GREENWOOD 96 search for neutrino oscillations at 18 m and 24 m from the reactor at Savannah River.

<sup>11</sup> DECLAIS 94 result based on integral measurement of neutrons only. Result is ratio of measured cross section to that expected in standard V-A theory. Replaced by ACHKAR 95.

<sup>12</sup> KWON 81 represents an analysis of a larger set of data from the same experiment as BOEHM 80.

### Atmospheric neutrinos

Neutrinos and antineutrinos produced in the atmosphere induce  $\mu$ -like and  $e$ -like events in underground detectors. The ratio of the numbers of the two kinds of events is defined as  $\mu/e$ . It has the advantage that systematic effects, such as flux uncertainty, tend to cancel, for both experimental and theoretical values of the ratio. The "ratio of the ratios" of experimental to theoretical  $\mu/e$ ,  $R(\mu/e)$ , or that of experimental to theoretical  $\mu/\text{total}$ ,  $R(\mu/\text{total})$  with  $\text{total} = \mu + e$ , is reported below. If the actual value is not unity, the value obtained in a given experiment may depend on the experimental conditions. In addition, the measured "up-down asymmetry" for  $\mu$  ( $N_{\text{up}}(\mu)/N_{\text{down}}(\mu)$ ) or  $e$  ( $N_{\text{up}}(e)/N_{\text{down}}(e)$ ) is reported. The expected "up-down asymmetry" is nearly unity if there is no neutrino oscillation.

### $R(\mu/e) = (\text{Measured Ratio } \mu/e) / (\text{Expected Ratio } \mu/e)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.658±0.016±0.035	<sup>1</sup> ASHIE 05	SKAM	sub-GeV
0.702±0.032±0.101 −0.030	<sup>2</sup> ASHIE 05	SKAM	multi-GeV
0.69 ±0.10 ±0.06	<sup>3</sup> SANCHEZ 03	SOU2	Calorimeter raw data
1.00 ±0.15 ±0.08	<sup>4</sup> FUKUDA 96B	KAMI	Water Cherenkov
0.60 +0.06 ±0.05 −0.05	<sup>5</sup> DAUM 95	FREJ	Calorimeter
0.57 +0.08 ±0.07 −0.07	<sup>6</sup> FUKUDA 94	KAMI	sub-GeV
	<sup>7</sup> FUKUDA 94	KAMI	multi-GeV
	<sup>8</sup> BECKER-SZ... 92B	IMB	Water Cherenkov

<sup>1</sup> ASHIE 05 results are based on an exposure of 92 kton yr during the complete Super-Kamiokande I running period. The analyzed data sample consists of fully-contained single-ring  $e$ -like events with  $0.1 \text{ GeV}/c < p_e$  and  $\mu$ -like events  $0.2 \text{ GeV}/c < p_\mu$ , both having a visible energy  $< 1.33 \text{ GeV}$ . These criteria match the definition used by FUKUDA 94.

<sup>2</sup> ASHIE 05 results are based on an exposure of 92 kton yr during the complete Super-Kamiokande I running period. The analyzed data sample consists of fully-contained single-ring events with visible energy  $> 1.33 \text{ GeV}$  and partially-contained events. All partially-contained events are classified as  $\mu$ -like.

<sup>3</sup> SANCHEZ 03 result is based on an exposure of 5.9 kton yr, and updates ALLISON 99 result. The analyzed data sample consists of fully-contained  $e$ -flavor and  $\mu$ -flavor events having lepton momentum  $> 0.3 \text{ GeV}/c$ .

<sup>4</sup> FUKUDA 96B studied neutron background in the atmospheric neutrino sample observed in the Kamiokande detector. No evidence for the background contamination was found.

<sup>5</sup> DAUM 95 results are based on an exposure of 2.0 kton yr which includes the data used by BERGER 90B. This ratio is for the contained and semicontained events. DAUM 95 also report  $R(\mu/e) = 0.99 \pm 0.13 \pm 0.08$  for the total neutrino induced data sample which includes upward going stopping muons and horizontal muons in addition to the contained and semicontained events.

<sup>6</sup> FUKUDA 94 result is based on an exposure of 7.7 kton yr and updates the HIRATA 92 result. The analyzed data sample consists of fully-contained  $e$ -like events with  $0.1 < p_e < 1.33 \text{ GeV}/c$  and fully-contained  $\mu$ -like events with  $0.2 < p_\mu < 1.5 \text{ GeV}/c$ .

<sup>7</sup> FUKUDA 94 analyzed the data sample consisting of fully contained events with visible energy  $> 1.33 \text{ GeV}$  and partially contained  $\mu$ -like events.

<sup>8</sup> BECKER-SZENDY 92B reports the fraction of nonshowing events (mostly muons from atmospheric neutrinos) as  $0.36 \pm 0.02 \pm 0.02$ , as compared with expected fraction  $0.51 \pm 0.01 \pm 0.05$ . After cutting the energy range to the Kamiokande limits, BEIER 92 finds  $R(\mu/e)$  very close to the Kamiokande value.

### $R(\nu_\mu) = (\text{Measured Flux of } \nu_\mu) / (\text{Expected Flux of } \nu_\mu)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.84±0.12	<sup>1</sup> ADAMSON 06	MINS	MINOS atmospheric
0.72±0.026±0.13	<sup>2</sup> AMBROSIO 01	MCRO	upward through-going
0.57±0.05 ±0.15	<sup>3</sup> AMBROSIO 00	MCRO	upgoing partially contained
0.71±0.05 ±0.19	<sup>4</sup> AMBROSIO 00	MCRO	downgoing partially contained + upgoing stopping
0.74±0.036±0.046	<sup>5</sup> AMBROSIO 98	MCRO	Streamer tubes
	<sup>6</sup> CASPER 91	IMB	Water Cherenkov
	<sup>7</sup> AGLIETTA 89	NUSX	
0.95±0.22	<sup>8</sup> BOLIEV 81		Baksan
0.62±0.17	CROUCH 78		Case Western/UCI

<sup>1</sup> ADAMSON 06 uses a measurement of 107 total neutrinos compared to an expected rate of 127 ± 13 without oscillations.

<sup>2</sup> AMBROSIO 01 result is based on the upward through-going muon tracks with  $E_\mu > 1 \text{ GeV}$ . The data came from three different detector configurations, but the statistics is largely dominated by the full detector run, from May 1994 to December 2000. The total live time, normalized to the full detector configuration, is 6.17 years. The first error is the statistical error, the second is the systematic error, dominated by the theoretical error in the predicted flux.

<sup>3</sup> AMBROSIO 00 result is based on the upgoing partially contained event sample. It came from 4.1 live years of data taking with the full detector, from April 1994 to February 1999. The average energy of atmospheric muon neutrinos corresponding to this sample is 4 GeV. The first error is statistical, the second is the systematic error, dominated by the 25% theoretical error in the rate (20% in the flux and 15% in the cross section, added in quadrature). Within statistics, the observed deficit is uniform over the zenith angle.

<sup>4</sup> AMBROSIO 00 result is based on the combined samples of downgoing partially contained events and upgoing stopping events. These two subsamples could not be distinguished due to the lack of timing information. The result came from 4.1 live years of data taking with the full detector, from April 1994 to February 1999. The average energy of atmospheric muon neutrinos corresponding to this sample is 4 GeV. The first error is statistical, the second is the systematic error, dominated by the 25% theoretical error in the rate (20% in the flux and 15% in the cross section, added in quadrature). Within statistics, the observed deficit is uniform over the zenith angle.

<sup>5</sup> AMBROSIO 98 result is for all nadir angles and updates AHLEN 95 result. The lower cutoff on the muon energy is 1 GeV. In addition to the statistical and systematic errors, there is a Monte Carlo flux error (theoretical error) of  $\approx 0.13$ . With a neutrino oscillation hypothesis, the fit either to the flux or zenith distribution independently yields  $\sin^2 2\theta = 1.0$  and  $\Delta(m^2) \sim$  a few times  $10^{-3} \text{ eV}^2$ . However, the fit to the observed zenith distribution gives a maximum probability for  $\chi^2$  of only 5% for the best oscillation hypothesis.

<sup>6</sup> CASPER 91 correlates showering/nonshowing signature of single-ring events with parent atmospheric-neutrino flavor. They find nonshowing ( $\approx \nu_\mu$  induced) fraction is  $0.41 \pm 0.03 \pm 0.02$ , as compared with expected  $0.51 \pm 0.05$  (syst).

<sup>7</sup> AGLIETTA 89 finds no evidence for any anomaly in the neutrino flux. They define  $\rho = (\text{measured number of } \nu_e \text{'s})/(\text{measured number of } \nu_\mu \text{'s})$ . They report  $\rho(\text{measured}) = \rho(\text{expected}) = 0.96 \pm 0.32$ .

<sup>8</sup> From this data BOLIEV 81 obtain the limit  $\Delta(m^2) \leq 6 \times 10^{-3} \text{ eV}^2$  for maximal mixing,  $\nu_\mu \not\leftrightarrow \nu_\mu$  type oscillation.

### $R(\mu/\text{total}) = (\text{Measured Ratio } \mu/\text{total}) / (\text{Expected Ratio } \mu/\text{total})$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.1+0.07 −0.12 ±0.11	<sup>1</sup> CLARK 97	IMB	multi-GeV

<sup>1</sup> CLARK 97 obtained this result by an analysis of fully contained and partially contained events in the IMB water-Cherenkov detector with visible energy  $> 0.95 \text{ GeV}$ .

### $N_{\text{up}}(\mu)/N_{\text{down}}(\mu)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			

# Lepton Particle Listings

## Neutrino Mixing

$0.71 \pm 0.06$	<sup>1</sup> ADAMSON	12B	MINS	contained-vertex muons
$0.551 \pm 0.035$ $-0.033 \pm 0.004$	<sup>2</sup> ASHIE	05	SKAM	multi-GeV

<sup>1</sup> ADAMSON 12B reports the atmospheric neutrino results obtained with MINOS far detector in 2,553 live days (an exposure of 37.9 kton-yr). This result is obtained with a sample of high resolution contained-vertex muons. The quoted error is statistical only.

<sup>2</sup> ASHIE 05 results are based on an exposure of 92 kton yr during the complete Super-Kamiokande I running period. The analyzed data sample consists of fully-contained single-ring  $\mu$ -like events with visible energy  $> 1.33$  GeV and partially-contained events. All partially-contained events are classified as  $\mu$ -like. Upward-going events are those with  $-1 < \cos(\text{zenith angle}) < -0.2$  and downward-going events are those with  $0.2 < \cos(\text{zenith angle}) < 1$ . The  $\mu$ -like up-down ratio for the multi-GeV data deviates from 1 (the expectation for no atmospheric  $\nu_\mu$  oscillations) by more than 12 standard deviations.

### $N_{\text{up}}(e)/N_{\text{down}}(e)$

VALUE	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.961 \pm 0.086$ $-0.079 \pm 0.016$	<sup>1</sup> ASHIE	05	SKAM	multi-GeV
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<sup>1</sup> ASHIE 05 results are based on an exposure of 92 kton yr during the complete Super-Kamiokande I running period. The analyzed data sample consists of fully-contained single-ring  $e$ -like events with visible energy  $> 1.33$  GeV. Upward-going events are those with  $-1 < \cos(\text{zenith angle}) < -0.2$  and downward-going events are those with  $0.2 < \cos(\text{zenith angle}) < 1$ . The  $e$ -like up-down ratio for the multi-GeV data is consistent with 1 (the expectation for no atmospheric  $\nu_e$  oscillations).

### $R(\text{up/down}; \mu) = (\text{Measured up/down}; \mu) / (\text{Expected up/down}; \mu)$

VALUE	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.62 \pm 0.05 \pm 0.02$	<sup>1</sup> ADAMSON	12B	MINS	contained-vertex muons
$0.62 \pm 0.19$ $-0.14 \pm 0.02$	<sup>2</sup> ADAMSON	06	MINS	atmospheric $\nu$ with far detector

<sup>1</sup> ADAMSON 12B reports the atmospheric neutrino results obtained with MINOS far detector in 2,553 live days (an exposure of 37.9 kton-yr). This result is obtained with a sample of high resolution contained-vertex muons. The expected ratio is calculated with no neutrino oscillation.

<sup>2</sup> ADAMSON 06 result is obtained with the MINOS far detector with an exposure of 4.54 kton yr. The expected ratio is calculated with no neutrino oscillation.

### $N(\mu^+)/N(\mu^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.46 \pm 0.05$ $-0.04$	<sup>1,2</sup> ADAMSON	12B	MINS	contained-vertex muons
$0.63 \pm 0.09$ $-0.08$	<sup>1,3</sup> ADAMSON	12B	MINS	$\nu$ -induced rock-muons

<sup>1</sup> ADAMSON 12B reports the atmospheric neutrino results obtained with MINOS far detector in 2,553 live days (an exposure of 37.9 kton-yr). The muon charge ratio  $N(\mu^+)/N(\mu^-)$  represents the  $\overline{\nu}_\mu/\nu_\mu$  ratio.

<sup>2</sup> This result is obtained with a charge-separated sample of high resolution contained-vertex muons. The quoted error is statistical only.

<sup>3</sup> This result is obtained with a charge-separated sample of high resolution neutrino-induced rock-muons. The quoted error is statistical only.

### $R(\mu^+/\mu^-) = (\text{Measured } N(\mu^+)/N(\mu^-)) / (\text{Expected } N(\mu^+)/N(\mu^-))$

VALUE	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.93 \pm 0.09 \pm 0.09$	<sup>1,2</sup> ADAMSON	12B	MINS	contained-vertex muons
$1.29 \pm 0.19$ $-0.17 \pm 0.16$	<sup>1,3</sup> ADAMSON	12B	MINS	$\nu$ -induced rock-muons
$1.03 \pm 0.08 \pm 0.08$	<sup>1,4</sup> ADAMSON	12B	MINS	contained
$1.39 \pm 0.35 \pm 0.08$ $-0.46 \pm 0.14$	<sup>5</sup> ADAMSON	07	MINS	Upward and horizontal $\mu$ with far detector
$0.96 \pm 0.38$ $-0.27 \pm 0.15$	<sup>6</sup> ADAMSON	06	MINS	atmospheric $\nu$ with far detector

<sup>1</sup> ADAMSON 12B reports the atmospheric neutrino results obtained with MINOS far detector in 2,553 live days (an exposure of 37.9 kton-yr). The muon charge ratio  $N(\mu^+)/N(\mu^-)$  represents the  $\overline{\nu}_\mu/\nu_\mu$  ratio. As far as the same oscillation parameters are used for  $\nu_s$  and  $\overline{\nu}_s$ , the expected  $\overline{\nu}_\mu/\nu_\mu$  ratio is almost entirely independent of any input oscillations.

<sup>2</sup> This result is obtained with a charge-separated sample of high resolution contained-vertex muons.

<sup>3</sup> This result is obtained with a charge-separated sample of high resolution neutrino-induced rock-muons.

<sup>4</sup> The charge-separated samples of high resolution contained-vertex muons and neutrino-induced rock-muons are combined to obtain this result which is consistent with unity.

<sup>5</sup> ADAMSON 07 result is obtained with the MINOS far detector in 854.24 live days, based on neutrino-induced upward-going and horizontal muons. This result is consistent with  $CPT$  conservation.

<sup>6</sup> ADAMSON 06 result is obtained with the MINOS far detector with an exposure of 4.54 kton yr, based on contained events. The expected ratio is calculated by assuming the same oscillation parameters for neutrinos and antineutrinos.

## Solar neutrinos

Solar neutrinos are produced by thermonuclear fusion reactions in the Sun. Radiochemical experiments measure particular combinations of fluxes from various neutrino-producing reactions, whereas water-Cherenkov experiments mainly measure a flux of neutrinos from decay of  $^8\text{B}$ . Solar neutrino fluxes are composed of all active neutrino species,  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ . In addition, some other mechanisms may cause antineutrino components in solar neutrino fluxes. Each measurement method is sensitive to a particular component or a combination of components of solar neutrino fluxes. For details, see Section 13.4 of Reviews, Tables, and Plots.

### $\nu_e$ Capture Rates from Radiochemical Experiments

1 SNU (Solar Neutrino Unit) =  $10^{-36}$  captures per atom per second.

VALUE (SNU)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$73.4 \pm 6.1 \pm 3.7$ $-6.0 \pm 4.1$	<sup>1</sup> KAETHER	10		GALX reanalysis
$67.6 \pm 4.0 \pm 3.2$	<sup>2</sup> KAETHER	10		GNO+GALX reanalysis combined
$65.4 \pm 3.1 \pm 2.6$ $-3.0 \pm 2.8$	<sup>3</sup> ABDURASHI...	09	SAGE	$^{71}\text{Ga} \rightarrow ^{71}\text{Ge}$
$62.9 \pm 5.5 \pm 2.5$ $-5.3$	<sup>4</sup> ALTMANN	05	GNO	$^{71}\text{Ga} \rightarrow ^{71}\text{Ge}$
$69.3 \pm 4.1 \pm 3.6$	<sup>5</sup> ALTMANN	05	GNO	GNO + GALX combined
$77.5 \pm 6.2 \pm 4.3$ $-4.7$	<sup>6</sup> HAMPEL	99	GALX	$^{71}\text{Ga} \rightarrow ^{71}\text{Ge}$
$2.56 \pm 0.16 \pm 0.16$	<sup>7</sup> CLEVELAND	98	HOME	$^{37}\text{Cl} \rightarrow ^{37}\text{Ar}$

<sup>1</sup> KAETHER 10 reports the reanalysis results of a complete GALLEX data (GALLEX I+II+III+IV, reported in HAMPEL 99) based on the event selection with a new pulse shape analysis, which provides a better background reduction than the rise time analysis adopted in HAMPEL 99.

<sup>2</sup> Combined result of GALLEX I+II+III+IV reanalysis and GNO I+II+III (ALTMANN 05).

<sup>3</sup> ABDURASHITOV 09 reports a combined analysis of 168 extractions of the SAGE solar neutrino experiment during the period January 1990 through December 2007, and updates the ABDURASHITOV 02 result. The data are consistent with the assumption that the solar neutrino production rate is constant in time. Note that a  $\sim 15\%$  systematic uncertainty in the overall normalization may be added to the ABDURASHITOV 09 result, because calibration experiments for gallium solar neutrino measurements using intense  $^{51}\text{Cr}$  (twice by GALLEX and once by SAGE) and  $^{37}\text{Ar}$  (by SAGE) result in an average ratio of  $0.87 \pm 0.05$  of the observed to calculated rates.

<sup>4</sup> ALTMANN 05 reports the complete result from the GNO solar neutrino experiment (GNO I+II+III), which is the successor project of GALLEX. Experimental technique of GNO is essentially the same as that of GALLEX. The run data cover the period 20 May 1998 through 9 April 2003.

<sup>5</sup> Combined result of GALLEX I+II+III+IV (HAMPEL 99) and GNO I+II+III.

<sup>6</sup> HAMPEL 99 report the combined result for GALLEX I+II+III+IV (65 runs in total), which update the HAMPEL 96 result. The GALLEX IV result (12 runs) is  $118.4 \pm 17.8 \pm 6.6$  SNU. (HAMPEL 99 discuss the consistency of partial results with the mean.) The GALLEX experimental program has been completed with these runs. The total run data cover the period 14 May 1991 through 23 January 1997. A total of 300  $^{71}\text{Ge}$  events were observed. Note that a  $\sim 15\%$  systematic uncertainty in the overall normalization may be added to the HAMPEL 99 result, because calibration experiments for gallium solar neutrino measurements using intense  $^{51}\text{Cr}$  (twice by GALLEX and once by SAGE) and  $^{37}\text{Ar}$  (by SAGE) result in an average ratio of  $0.87 \pm 0.05$  of the observed to calculated rates.

<sup>7</sup> CLEVELAND 98 is a detailed report of the  $^{37}\text{Cl}$  experiment at the Homestake Mine. The average solar neutrino-induced  $^{37}\text{Ar}$  production rate from 108 runs between 1970 and 1994 updates the DAVIS 89 result.

### $\Phi_{\text{ES}} (^8\text{B})$

$^8\text{B}$  solar-neutrino flux measured via  $\nu e$  elastic scattering. This process is sensitive to all active neutrino flavors, but with reduced sensitivity to  $\nu_\mu$ ,  $\nu_\tau$  due to the cross-section difference,  $\sigma(\nu_\mu, \tau e) \sim 0.16\sigma(\nu_e e)$ . If the  $^8\text{B}$  solar-neutrino flux involves nonelectron flavor active neutrinos, their contribution to the flux is  $\sim 0.16$  times of  $\nu_e$ .

VALUE ( $10^6 \text{ cm}^{-2}\text{s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.32 \pm 0.04 \pm 0.05$	<sup>1</sup> ABE	11	SKAM	SK-III average flux
$2.41 \pm 0.05 \pm 0.16$ $-0.15$	<sup>2</sup> ABE	11	SKAM	SK-II average flux
$2.38 \pm 0.02 \pm 0.08$	<sup>3</sup> ABE	11	SKAM	SK-I average flux
$2.77 \pm 0.26 \pm 0.32$	<sup>4</sup> ABE	11B	KLND	average flux
$2.4 \pm 0.4 \pm 0.1$	<sup>5</sup> BELLINI	10A	BORX	average flux
$1.77 \pm 0.24 \pm 0.09$ $-0.21 \pm 0.10$	<sup>6</sup> AHARMIM	08	SNO	Phase III
$2.38 \pm 0.05 \pm 0.16$ $-0.15$	<sup>7</sup> CRAVENS	08	SKAM	average flux
$2.35 \pm 0.02 \pm 0.08$	<sup>8</sup> HOSAKA	06	SKAM	average flux
$2.35 \pm 0.22 \pm 0.15$	<sup>9</sup> AHARMIM	05A	SNO	Salty $\text{D}_2\text{O}$ ; $^8\text{B}$ shape not constrained
$2.34 \pm 0.23 \pm 0.15$ $-0.14$	<sup>9</sup> AHARMIM	05A	SNO	Salty $\text{D}_2\text{O}$ ; $^8\text{B}$ shape constrained
$2.39 \pm 0.24 \pm 0.12$ $-0.23$	<sup>10</sup> AHMAD	02	SNO	average flux

$2.39 \pm 0.34^{+0.16}_{-0.14}$	11 AHMAD	01	SNO	average flux
$2.80 \pm 0.19 \pm 0.33$	12 FUKUDA	96	KAMI	average flux
$2.70 \pm 0.27$	12 FUKUDA	96	KAMI	day flux
$2.87^{+0.27}_{-0.26}$	12 FUKUDA	96	KAMI	night flux

<sup>1</sup> ABE 11 reports the Super-Kamiokande-III results for 548 live days from August 4, 2006 to August 18, 2008. The analysis threshold is 5.0 MeV, but the event sample in the 5.0–6.5 MeV total electron range has a total live time of 298 days.

<sup>2</sup> ABE 11 recalculated the Super-Kamiokande-II results using <sup>8</sup>B spectrum of WINTER 06A.

<sup>3</sup> ABE 11 recalculated the Super-Kamiokande-I results using <sup>8</sup>B spectrum of WINTER 06A.

<sup>4</sup> ABE 11B use a 123 kton-day exposure of the KamLAND liquid scintillation detector to measure the <sup>8</sup>B solar neutrino flux. They utilize  $\nu - e$  elastic scattering above a reconstructed-energy threshold of 5.5 MeV, corresponding to 5 MeV electron recoil energy. 299 electron recoil candidate events are reported, of which  $157 \pm 23.6$  are assigned to background.

<sup>5</sup> BELLINI 10A reports the Borexino result with 3 MeV energy threshold for scattered electrons. The data correspond to 345.3 live days with a target mass of 100 t, between July 15, 2007 and August 23, 2009.

<sup>6</sup> AHARMIM 08 reports the results from SNO Phase III measurement using an array of <sup>3</sup>He proportional counters to measure the rate of NC interactions in heavy water, over the period between November 27, 2004 and November 28, 2006, corresponding to 385.17 live days. A simultaneous fit was made for the number of NC events detected by the proportional counters and the numbers of NC, CC, and ES events detected by the PMTs, where the spectral distributions of the ES and CC events were not constrained to the <sup>8</sup>B shape.

<sup>7</sup> CRAVENS 08 reports the Super-Kamiokande-II results for 791 live days from December 2002 to October 2005. The photocathode coverage of the detector is 19% (reduced from 40% of that of Super-Kamiokande-I due to an accident in 2001). The analysis threshold for the average flux is 7 MeV.

<sup>8</sup> HOSAKA 06 reports the final results for 1496 live days with Super-Kamiokande-I between May 31, 1996 and July 15, 2001, and replace FUKUDA 02 results. The analysis threshold is 5 MeV except for the first 280 live days (6.5 MeV).

<sup>9</sup> AHARMIM 05A measurements were made with dissolved NaCl (0.195% by weight) in heavy water over the period between July 26, 2001 and August 28, 2003, corresponding to 391.4 live days, and update AHMED 04A. The CC, ES, and NC events were statistically separated. In one method, the <sup>8</sup>B energy spectrum was not constrained. In the other method, the constraint of an undistorted <sup>8</sup>B energy spectrum was added for comparison with AHMAD 02 results.

<sup>10</sup> AHMAD 02 reports the <sup>8</sup>B solar-neutrino flux measured via  $\nu e$  elastic scattering above the kinetic energy threshold of 5 MeV. The data correspond to 306.4 live days with SNO between November 2, 1999 and May 28, 2001, and updates AHMAD 01 results.

<sup>11</sup> AHMAD 01 reports the <sup>8</sup>B solar-neutrino flux measured via  $\nu e$  elastic scattering above the kinetic energy threshold of 6.75 MeV. The data correspond to 241 live days with SNO between November 2, 1999 and January 15, 2001.

<sup>12</sup> FUKUDA 96 results are for a total of 2079 live days with Kamiokande II and III from January 1987 through February 1995, covering the entire solar cycle 22, with threshold  $E_e > 9.3$  MeV (first 449 days),  $> 7.5$  MeV (middle 794 days), and  $> 7.0$  MeV (last 836 days). These results update the HIRATA 90 result for the average <sup>8</sup>B solar-neutrino flux and HIRATA 91 result for the day-night variation in the <sup>8</sup>B solar-neutrino flux. The total data sample was also analyzed for short-term variations: within experimental errors, no strong correlation of the solar-neutrino flux with the sunspot numbers was found.

### $\phi_{CC}$ (<sup>8</sup>B)

<sup>8</sup>B solar-neutrino flux measured with charged-current reaction which is sensitive exclusively to  $\nu_e$ .

VALUE ( $10^6 \text{ cm}^{-2}\text{s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.67^{+0.05+0.07}_{-0.04-0.08}$	1 AHARMIM	08	SNO Phase III
$1.68 \pm 0.06^{+0.08}_{-0.09}$	2 AHARMIM	05A	SNO Salty D <sub>2</sub> O; <sup>8</sup> B shape not const.
$1.72 \pm 0.05 \pm 0.11$	2 AHARMIM	05A	SNO Salty D <sub>2</sub> O; <sup>8</sup> B shape constrained
$1.76^{+0.06}_{-0.05} \pm 0.09$	3 AHMAD	02	SNO average flux
$1.75 \pm 0.07^{+0.12}_{-0.11} \pm 0.05$	4 AHMAD	01	SNO average flux

<sup>1</sup> AHARMIM 08 reports the results from SNO Phase III measurement using an array of <sup>3</sup>He proportional counters to measure the rate of NC interactions in heavy water, over the period between November 27, 2004 and November 28, 2006, corresponding to 385.17 live days. A simultaneous fit was made for the number of NC events detected by the proportional counters and the numbers of NC, CC, and ES events detected by the PMTs, where the spectral distributions of the ES and CC events were not constrained to the <sup>8</sup>B shape.

<sup>2</sup> AHARMIM 05A measurements were made with dissolved NaCl (0.195% by weight) in heavy water over the period between July 26, 2001 and August 28, 2003, corresponding to 391.4 live days, and update AHMED 04A. The CC, ES, and NC events were statistically separated. In one method, the <sup>8</sup>B energy spectrum was not constrained. In the other method, the constraint of an undistorted <sup>8</sup>B energy spectrum was added for comparison with AHMAD 02 results.

<sup>3</sup> AHMAD 02 reports the SNO result of the <sup>8</sup>B solar-neutrino flux measured with charged-current reaction on deuterium,  $\nu_e d \rightarrow ppe^-$ , above the kinetic energy threshold of 5 MeV. The data correspond to 306.4 live days with SNO between November 2, 1999 and May 28, 2001, and updates AHMAD 01 results. The complete description of the SNO Phase I data set is given in AHARMIM 07.

<sup>4</sup> AHMAD 01 reports the first SNO result of the <sup>8</sup>B solar-neutrino flux measured with the charged-current reaction on deuterium,  $\nu_e d \rightarrow ppe^-$ , above the kinetic energy threshold of 6.75 MeV. The data correspond to 241 live days with SNO between November 2, 1999 and January 15, 2001.

### $\phi_{NC}$ (<sup>8</sup>B)

<sup>8</sup>B solar neutrino flux measured with neutral-current reaction, which is equally sensitive to  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ .

VALUE ( $10^6 \text{ cm}^{-2}\text{s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$5.25 \pm 0.16^{+0.11}_{-0.13}$	1 AHARMIM	13	SNO All three phases combined
$5.140^{+0.160+0.132}_{-0.158-0.117}$	2 AHARMIM	10	SNO Phase I+II, low threshold
$5.54^{+0.33+0.36}_{-0.31-0.34}$	3 AHARMIM	08	SNO Phase III, prop. counter + PMT
$4.94 \pm 0.21^{+0.38}_{-0.34}$	4 AHARMIM	05A	SNO Salty D <sub>2</sub> O; <sup>8</sup> B shape not const.
$4.81 \pm 0.19^{+0.28}_{-0.27}$	4 AHARMIM	05A	SNO Salty D <sub>2</sub> O; <sup>8</sup> B shape constrained
$5.09^{+0.44+0.46}_{-0.43-0.43}$	5 AHMAD	02	SNO average flux; <sup>8</sup> B shape const.
$6.42 \pm 1.57^{+0.55}_{-0.58}$	5 AHMAD	02	SNO average flux; <sup>8</sup> B shape not const.

<sup>1</sup> AHARMIM 13 obtained this result from a combined analysis of the data from all three phases, SNO-I, II, and III. The measurement of the <sup>8</sup>B flux mostly comes from the NC signal, however, CC contribution is included in the fit.

<sup>2</sup> AHARMIM 10 reports this result from a joint analysis of SNO Phase I+II data with the "effective electron kinetic energy" threshold of 3.5 MeV. This result is obtained with a "binned-histogram unconstrained fit" where binned probability distribution functions of the neutrino signal observables were used without any model constraints on the shape of the neutrino spectrum.

<sup>3</sup> AHARMIM 08 reports the results from SNO Phase III measurement using an array of <sup>3</sup>He proportional counters to measure the rate of NC interactions in heavy water, over the period between November 27, 2004 and November 28, 2006, corresponding to 385.17 live days. A simultaneous fit was made for the number of NC events detected by the proportional counters and the numbers of NC, CC, and ES events detected by the PMTs, where the spectral distributions of the ES and CC events were not constrained to the <sup>8</sup>B shape.

<sup>4</sup> AHARMIM 05A measurements were made with dissolved NaCl (0.195% by weight) in heavy water over the period between July 26, 2001 and August 28, 2003, corresponding to 391.4 live days, and update AHMED 04A. The CC, ES, and NC events were statistically separated. In one method, the <sup>8</sup>B energy spectrum was not constrained. In the other method, the constraint of an undistorted <sup>8</sup>B energy spectrum was added for comparison with AHMAD 02 results.

<sup>5</sup> AHMAD 02 reports the first SNO result of the <sup>8</sup>B solar-neutrino flux measured with the neutral-current reaction on deuterium,  $\nu_\ell d \rightarrow np\nu_\ell$ , above the neutral-current reaction threshold of 2.2 MeV. The data correspond to 306.4 live days with SNO between November 2, 1999 and May 28, 2001. The complete description of the SNO Phase I data set is given in AHARMIM 07.

### $\phi_{\nu_\mu+\nu_\tau}$ (<sup>8</sup>B)

Nonelectron-flavor active neutrino component ( $\nu_\mu$  and  $\nu_\tau$ ) in the <sup>8</sup>B solar-neutrino flux.

VALUE ( $10^6 \text{ cm}^{-2}\text{s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.26 \pm 0.25^{+0.40}_{-0.35}$	1 AHARMIM	05A	SNO From $\phi_{NC}$ , $\phi_{CC}$ , and $\phi_{ES}$ ; <sup>8</sup> B shape not const.
$3.09 \pm 0.22^{+0.30}_{-0.27}$	1 AHARMIM	05A	SNO From $\phi_{NC}$ , $\phi_{CC}$ , and $\phi_{ES}$ ; <sup>8</sup> B shape constrained
$3.41 \pm 0.45^{+0.48}_{-0.45}$	2 AHMAD	02	SNO From $\phi_{NC}$ , $\phi_{CC}$ , and $\phi_{ES}$
$3.69 \pm 1.13$	3 AHMAD	01	SNO Derived from SNO+SuperKam, water Cherenkov

<sup>1</sup> AHARMIM 05A measurements were made with dissolved NaCl (0.195% by weight) in heavy water over the period between July 26, 2001 and August 28, 2003, corresponding to 391.4 live days, and update AHMED 04A. The CC, ES, and NC events were statistically separated. In one method, the <sup>8</sup>B energy spectrum was not constrained. In the other method, the constraint of an undistorted <sup>8</sup>B energy spectrum was added for comparison with AHMAD 02 results.

<sup>2</sup> AHMAD 02 deduced the nonelectron-flavor active neutrino component ( $\nu_\mu$  and  $\nu_\tau$ ) in the <sup>8</sup>B solar-neutrino flux, by combining the charged-current result, the  $\nu e$  elastic-scattering result and the neutral-current result. The complete description of the SNO Phase I data set is given in AHARMIM 07.

<sup>3</sup> AHMAD 01 deduced the nonelectron-flavor active neutrino component ( $\nu_\mu$  and  $\nu_\tau$ ) in the <sup>8</sup>B solar-neutrino flux, by combining the SNO charged-current result (AHMAD 01) and the Super-Kamiokande  $\nu e$  elastic-scattering result (FUKUDA 01).

### Total Flux of Active <sup>8</sup>B Solar Neutrinos

Total flux of active neutrinos ( $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ ).

VALUE ( $10^6 \text{ cm}^{-2}\text{s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			

# Lepton Particle Listings

## Neutrino Mixing

$5.25 \pm 0.16 \pm^{+0.11}_{-0.13}$	<sup>1</sup> AHARMIM	13	SNO	All three phases combined
$5.046 \pm^{+0.159}_{-0.152} \pm^{+0.107}_{-0.123}$	<sup>2</sup> AHARMIM	10	SNO	From $\phi_{NC}$ in Phase I+II, low threshold
$5.54 \pm^{+0.33}_{-0.31} \pm^{+0.36}_{-0.34}$	<sup>3</sup> AHARMIM	08	SNO	$\phi_{NC}$ in Phase III
$4.94 \pm 0.21 \pm^{+0.38}_{-0.34}$	<sup>4</sup> AHARMIM	05A	SNO	From $\phi_{NC}$ ; $^8\text{B}$ shape not const.
$4.81 \pm 0.19 \pm^{+0.28}_{-0.27}$	<sup>4</sup> AHARMIM	05A	SNO	From $\phi_{NC}$ ; $^8\text{B}$ shape constrained
$5.09 \pm^{+0.44}_{-0.43} \pm^{+0.46}_{-0.43}$	<sup>5</sup> AHMAD	02	SNO	Direct measurement from $\phi_{NC}$
$5.44 \pm 0.99$	<sup>6</sup> AHMAD	01		Derived from SNO+SuperKam, water Cherenkov

- <sup>1</sup> AHARMIM 13 obtained this result from a combined analysis of the data from all three phases, SNO-I, II, and III. The measurement of the  $^8\text{B}$  flux mostly comes from the NC signal, however, CC contribution is included in the fit.
- <sup>2</sup> AHARMIM 10 reports this result from a joint analysis of SNO Phase I+II data with the "effective electron kinetic energy" threshold of 3.5 MeV. This result is obtained with the assumption of unitarity, which relates the NC, CC, and ES rates. The data were fit with the free parameters directly describing the total  $^8\text{B}$  neutrino flux and the energy-dependent  $\nu_e$  survival probability.
- <sup>3</sup> AHARMIM 08 reports the results from SNO Phase III measurement using an array of  $^3\text{He}$  proportional counters to measure the rate of NC interactions in heavy water, over the period between November 27, 2004 and November 28, 2006, corresponding to 385.17 live days. A simultaneous fit was made for the number of NC events detected by the proportional counters and the numbers of NC, CC, and ES events detected by the PMTs, where the spectral distributions of the ES and CC events were not constrained to the  $^8\text{B}$  shape.
- <sup>4</sup> AHARMIM 05A measurements were made with dissolved NaCl (0.195% by weight) in heavy water over the period between July 26, 2001 and August 28, 2003, corresponding to 391.4 live days, and update AHMED 04A. The CC, ES, and NC events were statistically separated. In one method, the  $^8\text{B}$  energy spectrum was not constrained. In the other method, the constraint of an undistorted  $^8\text{B}$  energy spectrum was added for comparison with AHMAD 02 results.
- <sup>5</sup> AHMAD 02 determined the total flux of active  $^8\text{B}$  solar neutrinos by directly measuring the neutral-current reaction,  $\nu_e d \rightarrow np\nu_e$ , which is equally sensitive to  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ . The complete description of the SNO Phase I data set is given in AHARMIM 07.
- <sup>6</sup> AHMAD 01 deduced the total flux of active  $^8\text{B}$  solar neutrinos by combining the SNO charged-current result (AHMAD 01) and the Super-Kamiokande  $\nu e$  elastic-scattering result (FUKUDA 01).

### Day-Night Asymmetry ( $^8\text{B}$ )

$$A = (\phi_{\text{night}} - \phi_{\text{day}}) / \phi_{\text{average}}$$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.032 \pm 0.011 \pm 0.005$	<sup>1</sup> RENSHAW	14	SKAM Based on $\phi_{ES}$
$0.063 \pm 0.042 \pm 0.037$	<sup>2</sup> CRAVENS	08	SKAM Based on $\phi_{ES}$
$0.021 \pm 0.020 \pm^{+0.012}_{-0.013}$	<sup>3</sup> HOSAKA	06	SKAM Based on $\phi_{ES}$
$0.017 \pm 0.016 \pm^{+0.012}_{-0.013}$	<sup>4</sup> HOSAKA	06	SKAM Fitted in the LMA region
$-0.056 \pm 0.074 \pm 0.053$	<sup>5</sup> AHARMIM	05A	SNO From salty SNO $\phi_{CC}$
$-0.037 \pm 0.063 \pm 0.032$	<sup>5</sup> AHARMIM	05A	SNO From salty SNO $\phi_{CC}$ ; const. of no $\phi_{NC}$ asymmetry
$0.14 \pm 0.063 \pm^{+0.015}_{-0.014}$	<sup>6</sup> AHMAD	02B	SNO Derived from SNO $\phi_{CC}$
$0.07 \pm 0.049 \pm^{+0.013}_{-0.012}$	<sup>7</sup> AHMAD	02B	SNO Const. of no $\phi_{NC}$ asymmetry

- <sup>1</sup> RENSHAW 14 obtains this result by using the "amplitude fit" introduced in SMY 04. The data from the Super-Kamiokande(SK)-I, -II, and 1306 live days of the SK-IV measurements are used. The analysis threshold is recoil-electron kinetic energy of 4.5 MeV for SK-III, and SK-IV except for 250 live days in SK-III (6.0 MeV). The analysis threshold for SK-I and SK-II is the same as in the previous reports. (Note that in the previous SK solar-neutrino results, the analysis threshold is quoted as recoil-electron total energy.) This day-night asymmetry result is consistent with neutrino oscillations for  $4 \times 10^{-5} \text{ eV}^2 < \Delta m_{21}^2 < 7 \times 10^{-5} \text{ eV}^2$  and large mixing values of  $\theta_{12}$  at the 68% CL.
- <sup>2</sup> CRAVENS 08 reports the Super-Kamiokande-II results for 791 live days from December 2002 to October 2005. The photocathode coverage of the detector is 19% (reduced from 40% of that of Super-Kamiokande-I due to an accident in 2001). The analysis threshold for the day and night fluxes is 7.5 MeV except for the first 159 live days (8.0 MeV).
- <sup>3</sup> HOSAKA 06 reports the final results for 1496 live days with Super-Kamiokande-I between May 31, 1996 and July 15, 2001, and replace FUKUDA 02 results. The analysis threshold is 5 MeV except for the first 280 live days (6.5 MeV).
- <sup>4</sup> This result with reduced statistical uncertainty is obtained by assuming two-neutrino oscillations within the LMA (large mixing angle) region and by fitting the time variation of the solar neutrino flux measured via  $\nu_e$  elastic scattering to the variations expected from neutrino oscillations. For details, see SMY 04. There is an additional small systematic error of  $\pm 0.0004$  coming from uncertainty of oscillation parameters.
- <sup>5</sup> AHARMIM 05A measurements were made with dissolved NaCl (0.195% by weight) in heavy water over the period between July 26, 2001 and August 28, 2003, with 176.5 days of the live time recorded during the day and 214.9 days during the night. This result is obtained with the spectral distribution of the CC events not constrained to the  $^8\text{B}$  shape.
- <sup>6</sup> AHMAD 02B results are based on the charged-current interactions recorded between November 2, 1999 and May 28, 2001, with the day and night live times of 128.5 and 177.9 days, respectively. The complete description of the SNO Phase I data set is given in AHARMIM 07.

<sup>7</sup> AHMAD 02B results are derived from the charged-current interactions, neutral-current interactions, and  $\nu e$  elastic scattering, with the total flux of active neutrinos constrained to have no asymmetry. The data were recorded between November 2, 1999 and May 28, 2001, with the day and night live times of 128.5 and 177.9 days, respectively. The complete description of the SNO Phase I data set is given in AHARMIM 07.

### $\phi_{ES} (^8\text{Be})$

$^7\text{Be}$  solar-neutrino flux measured via  $\nu_e$  elastic scattering. This process is sensitive to all active neutrino flavors, but with reduced sensitivity to  $\nu_\mu$ ,  $\nu_\tau$  due to the cross-section difference,  $\sigma(\nu_{\mu,\tau} e) \sim 0.2 \sigma(\nu_e e)$ . If the  $^7\text{Be}$  solar-neutrino flux involves nonelectron flavor active neutrinos, their contribution to the flux is  $\sim 0.2$  times that of  $\nu_e$ .

VALUE ( $10^9 \text{ cm}^{-2} \text{ s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.26 \pm 0.52$	<sup>1</sup> GANDO	15	KLND average flux
$3.10 \pm 0.15$	<sup>2</sup> BELLINI	11A	BORX average flux

- <sup>1</sup> GANDO 15 uses 165.4 kton-day exposure of the KamLAND liquid scintillator detector to measure the 862 keV  $^7\text{Be}$  solar neutrino flux via  $\nu - e$  elastic scattering
- <sup>2</sup> BELLINI 11A reports the  $^7\text{Be}$  solar neutrino flux measured via  $\nu - e$  elastic scattering. The data correspond to 740.7 live days between May 16, 2007 and May 8, 2010, and also correspond to 153.6 ton-year fiducial exposure. BELLINI 11A measured the 862 keV  $^7\text{Be}$  solar neutrino flux, which is an 89.6% branch of the  $^7\text{Be}$  solar neutrino flux, to be  $(2.78 \pm 0.13) \times 10^9 \text{ cm}^{-2} \text{ s}^{-1}$ . Supercedes ARPESELLA 08A.

### $\phi_{ES} (p \bar{p})$

$p \bar{p}$  solar-neutrino flux measured via  $\nu_e$  elastic scattering. This process is sensitive to all active neutrino flavors, but with reduced sensitivity to  $\nu_\mu$ ,  $\nu_\tau$  due to the cross section difference,  $\sigma(\nu_{\mu,\tau} e) \sim 0.2 \sigma(\nu_e e)$ . If the  $p \bar{p}$  solar-neutrino flux involves non-electron flavor active neutrinos, their contribution to the flux is  $\sim 0.2$  times that of  $\nu_e$ .

VALUE ( $10^8 \text{ cm}^{-2} \text{ s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.0 \pm 0.2$	<sup>1</sup> BELLINI	12A	BORX average flux

- <sup>1</sup> BELLINI 12A reports 1.44 MeV  $p \bar{p}$  solar-neutrino flux measured via  $\nu_e$  elastic scattering. The data were collected between January 13, 2008 and May 9, 2010, corresponding to 20,4009 ton-day fiducial exposure. The listed flux value is calculated from the observed rate of  $p \bar{p}$  solar neutrino interactions in Borexino ( $3.1 \pm 0.6 \pm 0.3$  counts/(day-100 ton)) and the corresponding rate expected for no neutrino flavor oscillations ( $4.47 \pm 0.05$  counts/(day-100 ton)), using the SSM prediction for the  $p \bar{p}$  solar neutrino flux of  $(1.441 \pm 0.012) \times 10^8 \text{ cm}^{-2} \text{ s}^{-1}$ .

### $\phi_{ES} (\text{CNO})$

CNO solar-neutrino flux measured via  $\nu_e$  elastic scattering. This process is sensitive to all active neutrino flavors, but with reduced sensitivity to  $\nu_\mu$ ,  $\nu_\tau$  due to the cross section difference,  $\sigma(\nu_{\mu,\tau} e) \sim 0.2 \sigma(\nu_e e)$ . If the CNO solar-neutrino flux involves non-electron flavor active neutrinos, their contribution to the flux is  $\sim 0.2$  times that of  $\nu_e$ .

VALUE ( $10^8 \text{ cm}^{-2} \text{ s}^{-1}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 7.7$	90	<sup>1</sup> BELLINI	12A	BORX MSW-LMA solution assumed

- <sup>1</sup> BELLINI 12A reports an upper limit of the CNO solar neutrino flux measured via  $\nu_e$  elastic scattering. The data were collected between January 13, 2008 and May 9, 2010, corresponding to 20,409 ton-day fiducial exposure.

### $\phi_{ES} (pp)$

$pp$  solar-neutrino flux measured via  $\nu_e$  elastic scattering. This process is sensitive to all active neutrino flavors, but with reduced sensitivity to  $\nu_\mu$ ,  $\nu_\tau$  due to the cross section difference,  $\sigma(\nu_{\mu,\tau} e) \sim 0.3 \sigma(\nu_e e)$ . If the  $pp$  solar-neutrino flux involves nonelectron flavor active neutrinos, their contribution to the flux is  $\sim 0.3$  times of  $\nu_e$ .

VALUE ( $10^{10} \text{ cm}^{-2} \text{ s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$4.4 \pm 0.5$	<sup>1</sup> BELLINI	14A	BORX average flux

- <sup>1</sup> BELLINI 14A reports  $pp$  solar-neutrino flux measured via  $\nu_e$  elastic scattering. The data were collected between January 2012 and May 2013, corresponding to 408 days of data. The  $pp$  neutrino interaction rate in Borexino is measured to be  $144 \pm 13 \pm 10$  counts/(day-100 ton) by fitting the measured energy spectrum of events in the 165–590 keV recoil electron kinetic energy window with the expected signal + background spectrum. The listed flux value  $\phi_{ES}(pp)$  is calculated from the observed rate and the number of  $(3.307 \pm 0.003) \times 10^{31}$  electrons for 100 tons of the Borexino scintillator, and the  $\nu_e e$  integrated cross section over the  $pp$  neutrino spectrum,  $\sigma(\nu_e e) = 11.38 \times 10^{-46} \text{ cm}^2$ .

### $\phi_{CC} (pp)$

$pp$  solar-neutrino flux measured with charged-current reaction which is sensitive exclusively to  $\nu_e$ .

VALUE ( $10^{10} \text{ cm}^{-2} \text{ s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.38 \pm 0.47$	<sup>1</sup> ABDURASHI...	09	FIT Fit existing solar- $\nu$ data



<sup>1</sup> ABDURASHITOV 09 reports the  $pp$  solar-neutrino flux derived from the Ga solar-neutrino capture rate by subtracting contributions from  $^8\text{B}$ ,  $^7\text{Be}$ ,  $ppe$  and CNO solar-neutrino fluxes determined by other solar neutrino experiments as well as neutrino oscillation parameters determined from available world neutrino oscillation data.

### $\phi_{\text{ES}}$ (hep)

hep solar-neutrino flux measured via  $\nu e$  elastic scattering. This process is sensitive to all active neutrino flavors, but with reduced sensitivity to  $\nu_\mu$ ,  $\nu_\tau$  due to the cross-section difference,  $\sigma(\nu_{\mu,\tau}e) \sim 0.16\sigma(\nu_e e)$ . If the hep solar-neutrino flux involves nonelectron flavor active neutrinos, their contribution to the flux is  $\sim 0.16$  times of  $\nu_e$ .

VALUE ( $10^3 \text{ cm}^{-2} \text{ s}^{-1}$ )	CL%	DOCUMENT ID	TECN
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<73	90	<sup>1</sup> HOSAKA	06 SKAM
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<sup>1</sup> HOSAKA 06 result is obtained from the recoil electron energy window of 18–21 MeV, and updates FUKUDA 01 result.

### $\phi_{\overline{\nu}_e}$ ( $^8\text{B}$ )

Searches are made for electron antineutrino flux from the Sun. Flux limits listed here are derived relative to the BS05(OP) Standard Solar Model  $^8\text{B}$  solar neutrino flux ( $5.69 \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$ ), with an assumption that solar  $\overline{\nu}_e$ s follow an unoscillated  $^8\text{B}$  neutrino spectrum.

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.013	90	BELLINI	11 BORX	$E_{\overline{\nu}_e} > 1.8 \text{ MeV}$
<1.9	90	<sup>1</sup> BALATA	06 CNTR	$1.8 < E_{\overline{\nu}_e} < 20.0 \text{ MeV}$
<0.72	90	AHARMIM	04 SNO	$4.0 < E_{\overline{\nu}_e} < 14.8 \text{ MeV}$
<0.022	90	EGUCHI	04 KLND	$8.3 < E_{\overline{\nu}_e} < 14.8 \text{ MeV}$
<0.7	90	GANDO	03 SKAM	$8.0 < E_{\overline{\nu}_e} < 20.0 \text{ MeV}$
<1.7	90	AGLIETTA	96 LSD	$7 < E_{\overline{\nu}_e} < 17 \text{ MeV}$

<sup>1</sup> BALATA 06 obtained this result from the search for  $\overline{\nu}_e$  interactions with Counting Test Facility (the prototype of the Borexino detector).

## (B) Three-neutrino mixing parameters

## INTRODUCTION TO THREE-NEUTRINO MIXING PARAMETERS LISTINGS

Updated November 2015 by M. Goodman (ANL).

**Introduction and Notation:** With the exception of possible short-baseline anomalies (such as LSND), current accelerator, reactor, solar and atmospheric neutrino data can be described within the framework of a  $3 \times 3$  mixing matrix between the flavor eigenstates  $\nu_e$ ,  $\nu_\mu$  and  $\nu_\tau$  and mass eigenstates  $\nu_1$ ,  $\nu_2$  and  $\nu_3$ . (See equation 14.6 of the review “Neutrino Mass, Mixing and Oscillations” by K. Nakamura and S.T. Petcov.) Whether or not this is the ultimately correct framework, it is currently widely used to parametrize neutrino mixing data and to plan new experiments.

The mass differences are called  $\Delta m_{21}^2 \equiv m_2^2 - m_1^2$  and  $\Delta m_{32}^2 \equiv m_3^2 - m_2^2$ . In these listings, we assume

$$\Delta m_{32}^2 \sim \Delta m_{31}^2 \quad (1)$$

even though the experimental error is comparable to the difference  $\Delta m_{31}^2 - \Delta m_{32}^2 = \Delta m_{21}^2$ . The measurements made by  $\nu_\mu$  disappearance at accelerators and by  $\nu_e$  disappearance at reactors are slightly different mixtures of  $\Delta m_{32}^2$  and  $\Delta m_{31}^2$ . The angles are labeled  $\theta_{12}$ ,  $\theta_{23}$  and  $\theta_{13}$ . The CP violating phase is called  $\delta$ . The familiar two neutrino form for oscillations is

$$P(\nu_a \rightarrow \nu_b; a \neq b) = \sin^2(2\theta) \sin^2(\Delta m^2 L/4E). \quad (2)$$

Despite the fact that the mixing angles have been measured to be much larger than in the quark sector, the two neutrino form is often a very good approximation and is used in many situations.

The angles appear in the equations below in many forms. They most often appear as  $\sin^2(2\theta)$ . The listings currently now use  $\sin^2(\theta)$  because this distinguishes whether  $\theta_{23}$  is larger or smaller than  $45^\circ$ .

**Accelerator neutrino experiments:** Ignoring  $\Delta m_{21}^2$ , CP violation, and matter effects, the equations for the probability of appearance in an accelerator oscillation experiment are:

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2(2\theta_{23}) \cos^4(\theta_{13}) \sin^2(\Delta m_{32}^2 L/4E) \quad (3)$$

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2(2\theta_{13}) \sin^2(\theta_{23}) \sin^2(\Delta m_{32}^2 L/4E) \quad (4)$$

$$P(\nu_e \rightarrow \nu_\mu) = \sin^2(2\theta_{13}) \sin^2(\theta_{23}) \sin^2(\Delta m_{32}^2 L/4E) \quad (5)$$

$$P(\nu_e \rightarrow \nu_\tau) = \sin^2(2\theta_{13}) \cos^2(\theta_{23}) \sin^2(\Delta m_{32}^2 L/4E). \quad (6)$$

Current and future long-baseline accelerator experiments are studying non-zero  $\theta_{13}$  through  $P(\nu_\mu \rightarrow \nu_e)$ . Including the CP terms and low mass scale, the equation for neutrino oscillation in vacuum is:

$$\begin{aligned} P(\nu_\mu \rightarrow \nu_e) &= P1 + P2 + P3 + P4 \\ P1 &= \sin^2(\theta_{23}) \sin^2(2\theta_{13}) \sin^2(\Delta m_{32}^2 L/4E) \\ P2 &= \cos^2(\theta_{23}) \sin^2(2\theta_{13}) \sin^2(\Delta m_{21}^2 L/4E) \\ P3 &= -/+ J \sin(\delta) \sin(\Delta m_{32}^2 L/4E) \\ P4 &= J \cos(\delta) \cos(\Delta m_{32}^2 L/4E) \end{aligned} \quad (7)$$

where

$$\begin{aligned} J &= \cos(\theta_{13}) \sin(2\theta_{12}) \sin(2\theta_{13}) \sin(2\theta_{23}) \times \\ &\quad \sin(\Delta m_{32}^2 L/4E) \sin(\Delta m_{21}^2 L/4E) \end{aligned} \quad (8)$$

and the sign in P3 is negative for neutrinos and positive for anti-neutrinos respectively. For most new long-baseline accelerator experiments, P2 can safely be neglected but the other three terms can all be large. Also, depending on the distance and the mass hierarchy, matter effects will need to be included.

**Reactor neutrino experiments:** Nuclear reactors are prolific sources of  $\bar{\nu}_e$  with an energy near 4 MeV. The oscillation probability can be expressed

$$\begin{aligned} P(\bar{\nu}_e \rightarrow \bar{\nu}_e) &= 1 - \cos^4(\theta_{13}) \sin^2(2\theta_{12}) \sin^2(\Delta m_{21}^2 L/4E) \\ &\quad - \cos^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta m_{31}^2 L/4E) \\ &\quad - \sin^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta m_{32}^2 L/4E) \end{aligned} \quad (9)$$

not using the approximation in Eq. (1). For short distances ( $L < 5 \text{ km}$ ) we can ignore the second term on the right and can reimpose approximation Eq. (1). This takes the familiar two neutrino form with  $\theta_{13}$  and  $\Delta m_{32}^2$ :

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2(2\theta_{13}) \sin^2(\Delta m_{32}^2 L/4E). \quad (10)$$

**Solar and Atmospheric neutrino experiments:** Solar neutrino experiments are sensitive to  $\nu_e$  disappearance and have

# Lepton Particle Listings

## Neutrino Mixing

allowed the measurement of  $\theta_{12}$  and  $\Delta m_{21}^2$ . They are also sensitive to  $\theta_{13}$ . We identify  $\Delta m_{\odot}^2 = \Delta m_{21}^2$  and  $\theta_{\odot} = \theta_{12}$ .

Atmospheric neutrino experiments are primarily sensitive to  $\nu_{\mu}$  disappearance through  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillations, and have allowed the measurement of  $\theta_{23}$  and  $\Delta m_{32}^2$ . We identify  $\Delta m_A^2 = \Delta m_{32}^2$  and  $\theta_A = \theta_{23}$ . Despite the large  $\nu_e$  component of the atmospheric neutrino flux, it is difficult to measure  $\Delta m_{21}^2$  effects. This is because of a cancellation between  $\nu_{\mu} \rightarrow \nu_e$  and  $\nu_e \rightarrow \nu_{\mu}$  together with the fact that the ratio of  $\nu_{\mu}$  and  $\nu_e$  atmospheric fluxes, which arise from sequential  $\pi$  and  $\mu$  decay, is near 2.

**Oscillation Parameter Listings:** In Section (B) we encode the three mixing angles  $\theta_{12}$ ,  $\theta_{23}$ ,  $\theta_{13}$  and two mass squared differences  $\Delta m_{21}^2$  and  $\Delta m_{32}^2$ . Our knowledge of  $\theta_{12}$  and  $\Delta m_{21}^2$  comes from the KamLAND reactor neutrino experiment together with solar neutrino experiments. Our knowledge of  $\theta_{23}$  and  $\Delta m_{32}^2$  comes from atmospheric, reactor and long-baseline accelerator neutrino experiments. For the earlier experiments, we identified the large mass splitting as  $\Delta m_{32}^2$ . Now that  $\sigma(\Delta m_{32}^2) \approx \Delta m_{21}^2$ , some experiments report separate values for the two hierarchies. Results on  $\theta_{13}$  come from reactor antineutrino disappearance experiments. There are also results from long-baseline accelerator experiments looking for  $\nu_e$  appearance. The interpretation of both kinds of results depends on  $\Delta m_{32}^2$ , and the accelerator results also depend on the mass hierarchy,  $\theta_{23}$  and the CP violating phase  $\delta$ .

Accelerator and atmospheric experiments are beginning to have some sensitivity to the CP violation phase  $\delta$  through Eq. (7). Note that P3 depends on the sign of  $\Delta m_{32}^2$  so the sensitivity depends on the mass hierarchy. For non-maximal  $\theta_{23}$  mixing, it also depends on the octant of  $\theta_{23}$ , i.e. whether  $\theta_{23} > \pi/4$  or  $\theta_{23} < \pi/4$ .

### $\sin^2(\theta_{12})$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.304^{+0.014}_{-0.013}$	<sup>1</sup> GANDO	13 FIT	KamLAND + global solar + SBL + accelerator: $3\nu$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.323 \pm 0.016$	<sup>2</sup> FORERO	14 FIT	$3\nu$
$0.304^{+0.013}_{-0.012}$	<sup>3</sup> GONZALEZ-G.	14 FIT	Either mass ordering; global fit
$0.299^{+0.014}_{-0.014}$	<sup>4,5</sup> AHARMIM	13 FIT	global solar: $2\nu$
$0.307^{+0.016}_{-0.013}$	<sup>5,6</sup> AHARMIM	13 FIT	global solar: $3\nu$
$0.304^{+0.022}_{-0.018}$	<sup>5,7</sup> AHARMIM	13 FIT	KamLAND + global solar: $3\nu$
$0.304^{+0.014}_{-0.013}$	<sup>8</sup> GANDO	13 FIT	KamLAND + global solar: $3\nu$
$0.325^{+0.039}_{-0.039}$	<sup>9</sup> GANDO	13 FIT	KamLAND: $3\nu$
$0.30^{+0.02}_{-0.01}$	<sup>10</sup> ABE	11 FIT	KamLAND + global solar: $2\nu$
$0.30^{+0.02}_{-0.01}$	<sup>11</sup> ABE	11 FIT	global solar: $2\nu$
$0.31^{+0.03}_{-0.02}$	<sup>12</sup> ABE	11 FIT	KamLAND + global solar: $3\nu$
$0.31^{+0.03}_{-0.03}$	<sup>13</sup> ABE	11 FIT	global solar: $3\nu$
$0.314^{+0.015}_{-0.012}$	<sup>14</sup> BELLINI	11A FIT	KamLAND + global solar: $2\nu$
$0.319^{+0.017}_{-0.015}$	<sup>15</sup> BELLINI	11A FIT	global solar: $2\nu$
$0.311^{+0.016}_{-0.016}$	<sup>16</sup> GANDO	11 FIT	KamLAND + solar: $3\nu$
$0.304^{+0.046}_{-0.042}$	<sup>17</sup> GANDO	11 FIT	KamLAND: $3\nu$
$0.314^{+0.018}_{-0.014}$	<sup>18,19</sup> AHARMIM	10 FIT	KamLAND + global solar: $2\nu$

$0.314^{+0.017}_{-0.020}$	18,20 AHARMIM	10 FIT	global solar: $2\nu$
$0.319^{+0.019}_{-0.016}$	18,21 AHARMIM	10 FIT	KamLAND + global solar: $3\nu$
$0.319^{+0.023}_{-0.024}$	18,22 AHARMIM	10 FIT	global solar: $3\nu$
$0.36^{+0.05}_{-0.04}$	<sup>23</sup> ABE	08A FIT	KamLAND
$0.32 \pm 0.03$	<sup>24</sup> ABE	08A FIT	KamLAND + global fit
$0.32 \pm 0.02$	<sup>25</sup> AHARMIM	08 FIT	KamLAND + global solar
$0.31^{+0.04}_{-0.04}$	<sup>26</sup> HOSAKA	06 FIT	KamLAND + global solar
$0.31^{+0.04}_{-0.03}$	<sup>27</sup> HOSAKA	06 FIT	SKAM+SNO+KamLAND
$0.31^{+0.03}_{-0.04}$	<sup>28</sup> HOSAKA	06 FIT	SKAM+SNO
$0.31^{+0.02}_{-0.03}$	<sup>29</sup> AHARMIM	05A FIT	KamLAND + global solar
$0.25-0.39$	<sup>30</sup> AHARMIM	05A FIT	global solar
$0.29 \pm 0.03$	<sup>31</sup> ARAKI	05 FIT	KamLAND + global solar
$0.29^{+0.03}_{-0.02}$	<sup>32</sup> AHMED	04A FIT	KamLAND + global solar
$0.23-0.37$	<sup>33</sup> AHMED	04A FIT	global solar
$0.31^{+0.04}_{-0.04}$	<sup>34</sup> SMY	04 FIT	KamLAND + global solar
$0.29^{+0.04}_{-0.04}$	<sup>35</sup> SMY	04 FIT	global solar
$0.32^{+0.06}_{-0.05}$	<sup>36</sup> SMY	04 FIT	SKAM + SNO
$0.19-0.33$	<sup>37</sup> AHMAD	02B FIT	global solar
$0.19-0.39$	<sup>38</sup> FUKUDA	02 FIT	global solar

<sup>1</sup> GANDO 13 obtained this result by a three-neutrino oscillation analysis using KamLAND, global solar neutrino, short-baseline (SBL) reactor, and accelerator data, assuming CPT invariance. Supersedes GANDO 11.

<sup>2</sup> FORERO 14 performs a global fit to neutrino oscillations using solar, reactor, long-baseline accelerator, and atmospheric neutrino data.

<sup>3</sup> GONZALEZ-GARCIA 14 result comes from a frequentist global fit. The corresponding Bayesian global fit to the same data results are reported in BERGSTROM 15 as  $0.304^{+0.013}_{-0.012}$  for normal and  $0.305^{+0.012}_{-0.013}$  for inverted mass ordering.

<sup>4</sup> AHARMIM 13 obtained this result by a two-neutrino oscillation analysis using global solar neutrino data.

<sup>5</sup> AHARMIM 13 global solar neutrino data include SNO's all-phases-combined analysis results on the total active  $^8\text{B}$  neutrino flux and energy-dependent  $\nu_e$  survival probability parameters, measurements of Cl (CLEVELAND 98), Ga (ABDURASHITOV 09 which contains combined analysis with GNO (ALTMANN 05 and Ph.D. thesis of F. Kaether)), and  $^7\text{Be}$  (BELLINI 11A) rates, and  $^8\text{B}$  solar-neutrino recoil electron measurements of SK-I (HOSAKA 06) zenith, SK-II (CRAVENS 08) and SK-III (ABE 11) day/night spectra, and Borexino (BELLINI 10A) spectra.

<sup>6</sup> AHARMIM 13 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.45 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data.

<sup>7</sup> AHARMIM 13 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.45 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino and KamLAND (GANDO 11) data. CPT invariance is assumed.

<sup>8</sup> GANDO 13 obtained this result by a three-neutrino oscillation analysis using KamLAND and global solar neutrino data, assuming CPT invariance. Supersedes GANDO 11.

<sup>9</sup> GANDO 13 obtained this result by a three-neutrino oscillation analysis using KamLAND data. Supersedes GANDO 11.

<sup>10</sup> ABE 11 obtained this result by a two-neutrino oscillation analysis using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, SAGE, and KamLAND data. CPT invariance is assumed.

<sup>11</sup> ABE 11 obtained this result by a two-neutrino oscillation analysis using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, and SAGE data.

<sup>12</sup> ABE 11 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.4 \times 10^{-3} \text{ eV}^2$ , using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, SAGE, and KamLAND data. The normal neutrino mass ordering and CPT invariance are assumed.

<sup>13</sup> ABE 11 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.4 \times 10^{-3} \text{ eV}^2$ , using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, and GALLEX/GNO data. The normal neutrino mass ordering is assumed.

<sup>14</sup> BELLINI 11A obtained this result by a two-neutrino oscillation analysis using KamLAND, Homestake, SAGE, Gallex, GNO, Kamiokande, Super-Kamiokande, SNO, and Borexino (BELLINI 11A) data and the SSM flux prediction in SERENELLI 11 (Astrophysical Journal **743** 24 (2011)) with the exception that the  $^8\text{B}$  flux was left free. CPT invariance is assumed.

<sup>15</sup> BELLINI 11A obtained this result by a two-neutrino oscillation analysis using Homestake, SAGE, Gallex, GNO, Kamiokande, Super-Kamiokande, SNO, and Borexino (BELLINI 11A) data and the SSM flux prediction in SERENELLI 11 (Astrophysical Journal **743** 24 (2011)) with the exception that the  $^8\text{B}$  flux was left free.

<sup>16</sup> GANDO 11 obtain this result with three-neutrino fit using the KamLAND + solar data. Superseded by GANDO 13.

<sup>17</sup> GANDO 11 obtain this result with three-neutrino fit using the KamLAND data only. Superseded by GANDO 13.

<sup>18</sup> AHARMIM 10 global solar neutrino data include SNO's low-energy-threshold analysis survival probability day/night curves, SNO Phase III integral rates (AHARMIM 08), Cl (CLEVELAND 98), SAGE (ABDURASHITOV 09), Gallex/GNO (HAMPEL 99, ALTMANN 05), Borexino (ARPESELLA 08A), SK-I zenith (HOSAKA 06), and SK-II day/night spectra (CRAVENS 08).

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# Lepton Particle Listings

## Neutrino Mixing

- <sup>19</sup>AHARMIM 10 obtained this result by a two-neutrino oscillation analysis using global solar neutrino data and KamLAND data (ABE 08A). *CPT* invariance is assumed.
- <sup>20</sup>AHARMIM 10 obtained this result by a two-neutrino oscillation analysis using global solar neutrino data.
- <sup>21</sup>AHARMIM 10 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.3 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data and KamLAND data (ABE 08A). *CPT* invariance is assumed.
- <sup>22</sup>AHARMIM 10 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.3 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data.
- <sup>23</sup>ABE 08A obtained this result by a rate + shape + time combined geoneutrino and reactor two-neutrino fit for  $\Delta m_{21}^2$  and  $\tan^2 \theta_{12}$ , using KamLAND data only. Superseded by GANDO 11.
- <sup>24</sup>ABE 08A obtained this result by means of a two-neutrino fit using KamLAND, Homestake, SAGE, GALLEX, GNO, SK (zenith angle and E-spectrum), the SNO  $\chi^2$ -map, and solar flux data. *CPT* invariance is assumed. Superseded by GANDO 11.
- <sup>25</sup>The result given by AHARMIM 08 is  $\theta = (34.4^{+1.3}_{-1.2})^\circ$ . This result is obtained by a two-neutrino oscillation analysis using solar neutrino data including those of Borexino (ARPESELLA 08A) and Super-Kamiokande-I (HOSAKA 06), and KamLAND data (ABE 08A). *CPT* invariance is assumed.
- <sup>26</sup>HOSAKA 06 obtained this result by a two-neutrino oscillation analysis using SK  $\nu_e$  data, CC data from other solar neutrino experiments, and KamLAND data (ARAKI 05). *CPT* invariance is assumed.
- <sup>27</sup>HOSAKA 06 obtained this result by a two-neutrino oscillation analysis using the data from Super-Kamiokande, SNO (AHMAD 02 and AHMAD 02b), and KamLAND (ARAKI 05) experiments. *CPT* invariance is assumed.
- <sup>28</sup>HOSAKA 06 obtained this result by a two-neutrino oscillation analysis using the Super-Kamiokande and SNO (AHMAD 02 and AHMAD 02b) solar neutrino data.
- <sup>29</sup>The result given by AHARMIM 05A is  $\theta = (33.9 \pm 1.6)^\circ$ . This result is obtained by a two-neutrino oscillation analysis using SNO pure deuteron and salt phase data, SK  $\nu_e$  data, CI and Ga CC data, and KamLAND data (ARAKI 05). *CPT* invariance is assumed. AHARMIM 05A also quotes  $\theta = (33.9^{+2.4}_{-2.2})^\circ$  as the error enveloping the 68% CL two-dimensional region. This translates into  $\sin^2 2\theta = 0.86^{+0.05}_{-0.06}$ .
- <sup>30</sup>AHARMIM 05A obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in figure 35a of AHARMIM 05A. AHARMIM 05A also quotes  $\tan^2 \theta = 0.45^{+0.09}_{-0.08}$  as the error enveloping the 68% CL two-dimensional region. This translates into  $\sin^2 2\theta = 0.86^{+0.05}_{-0.07}$ .
- <sup>31</sup>ARAKI 05 obtained this result by a two-neutrino oscillation analysis using KamLAND and solar neutrino data. *CPT* invariance is assumed. The  $1\sigma$  error shown here is translated from the number provided by the KamLAND collaboration,  $\tan^2 \theta = 0.40^{+0.07}_{-0.05}$ . The corresponding number quoted in ARAKI 05 is  $\tan^2 \theta = 0.40^{+0.10}_{-0.07}$  ( $\sin^2 2\theta = 0.82 \pm 0.07$ ), which envelops the 68% CL two-dimensional region.
- <sup>32</sup>The result given by AHMED 04A is  $\theta = (32.5^{+1.7}_{-1.6})^\circ$ . This result is obtained by a two-neutrino oscillation analysis using solar neutrino and KamLAND data (EGUCHI 03). *CPT* invariance is assumed. AHMED 04A also quotes  $\theta = (32.5^{+2.4}_{-2.3})^\circ$  as the error enveloping the 68% CL two-dimensional region. This translates into  $\sin^2 2\theta = 0.82 \pm 0.06$ .
- <sup>33</sup>AHMED 04A obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in Fig. 5(a) of AHMED 04A. The best-fit point is  $\Delta(m^2) = 6.5 \times 10^{-5} \text{ eV}^2$ ,  $\tan^2 \theta = 0.40$  ( $\sin^2 2\theta = 0.82$ ).
- <sup>34</sup>The result given by SMY 04 is  $\tan^2 \theta = 0.44 \pm 0.08$ . This result is obtained by a two-neutrino oscillation analysis using solar neutrino and KamLAND data (IANNI 03). *CPT* invariance is assumed.
- <sup>35</sup>SMY 04 obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The  $1\sigma$  errors are read from Fig. 6(a) of SMY 04.
- <sup>36</sup>SMY 04 obtained this result by a two-neutrino oscillation analysis using the Super-Kamiokande and SNO (AHMAD 02 and AHMAD 02b) solar neutrino data. The  $1\sigma$  errors are read from Fig. 6(a) of SMY 04.
- <sup>37</sup>AHMAD 02b obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in Fig. 4(b) of AHMAD 02b. The best fit point is  $\Delta(m^2) = 5.0 \times 10^{-5} \text{ eV}^2$  and  $\tan \theta = 0.34$  ( $\sin^2 2\theta = 0.76$ ).
- <sup>38</sup>FUKUDA 02 obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in Fig. 4 of FUKUDA 02. The best fit point is  $\Delta(m^2) = 6.9 \times 10^{-5} \text{ eV}^2$  and  $\tan^2 \theta = 0.38$  ( $\sin^2 2\theta = 0.80$ ).

### $\Delta m_{21}^2$

VALUE ( $10^{-5} \text{ eV}^2$ )	DOCUMENT ID	TECN	COMMENT
<b><math>7.53 \pm 0.18</math></b>	<sup>1</sup> GANDO	13	FIT KamLAND + global solar + SBL + accelerator: $3\nu$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$7.6^{+0.19}_{-0.18}$	<sup>2</sup> FORERO	14	FIT $3\nu$
$7.50^{+0.19}_{-0.17}$	<sup>3</sup> GONZALEZ-G.	14	FIT Either mass ordering; global fit
$5.13^{+1.29}_{-0.96}$	<sup>4,5</sup> AHARMIM	13	FIT global solar: $2\nu$
$5.13^{+1.49}_{-0.98}$	<sup>5,6</sup> AHARMIM	13	FIT global solar: $3\nu$
$7.46^{+0.20}_{-0.19}$	<sup>5,7</sup> AHARMIM	13	FIT KamLAND + global solar: $3\nu$

$7.53^{+0.19}_{-0.18}$	<sup>8</sup> GANDO	13	FIT KamLAND + global solar: $3\nu$
$7.54^{+0.19}_{-0.18}$	<sup>9</sup> GANDO	13	FIT KamLAND: $3\nu$
$7.6 \pm 0.2$	<sup>10</sup> ABE	11	FIT KamLAND + global solar: $2\nu$
$6.2^{+1.1}_{-1.9}$	<sup>11</sup> ABE	11	FIT global solar: $2\nu$
$7.7 \pm 0.3$	<sup>12</sup> ABE	11	FIT KamLAND + global solar: $3\nu$
$6.0^{+2.2}_{-2.5}$	<sup>13</sup> ABE	11	FIT global solar: $3\nu$
$7.50^{+0.16}_{-0.24}$	<sup>14</sup> BELLINI	11A	FIT KamLAND + global solar: $2\nu$
$5.2^{+1.5}_{-0.9}$	<sup>15</sup> BELLINI	11A	FIT global solar: $2\nu$
$7.50^{+0.19}_{-0.20}$	<sup>16</sup> GANDO	11	FIT KamLAND + solar: $3\nu$
$7.49 \pm 0.20$	<sup>17</sup> GANDO	11	FIT KamLAND: $3\nu$
$7.50^{+0.20}_{-0.21}$	<sup>18,19</sup> AHARMIM	10	FIT KamLAND + global solar: $2\nu$
$5.89^{+2.13}_{-2.16}$	<sup>18,20</sup> AHARMIM	10	FIT global solar: $2\nu$
$7.59 \pm 0.21$	<sup>18,21</sup> AHARMIM	10	FIT KamLAND + global solar: $3\nu$
$6.31^{+2.49}_{-2.58}$	<sup>18,22</sup> AHARMIM	10	FIT global solar: $3\nu$
$7.58^{+0.14}_{-0.13} \pm 0.15$	<sup>23</sup> ABE	08A	FIT KamLAND
$7.59 \pm 0.21$	<sup>24</sup> ABE	08A	FIT KamLAND + global solar
$7.59^{+0.19}_{-0.21}$	<sup>25</sup> AHARMIM	08	FIT KamLAND + global solar
$8.0 \pm 0.3$	<sup>26</sup> HOSAKA	06	FIT KamLAND + global solar
$8.0 \pm 0.3$	<sup>27</sup> HOSAKA	06	FIT SKAM+SNO+KamLAND
$6.3^{+3.7}_{-1.5}$	<sup>28</sup> HOSAKA	06	FIT SKAM+SNO
5–12	<sup>29</sup> HOSAKA	06	FIT SKAM day/night in the LMA region
$8.0^{+0.4}_{-0.3}$	<sup>30</sup> AHARMIM	05A	FIT KamLAND + global solar LMA
3.3–14.4	<sup>31</sup> AHARMIM	05A	FIT global solar
$7.9^{+0.4}_{-0.3}$	<sup>32</sup> ARAKI	05	FIT KamLAND + global solar
$7.1^{+1.0}_{-0.3}$	<sup>33</sup> AHMED	04A	FIT KamLAND + global solar
3.2–13.7	<sup>34</sup> AHMED	04A	FIT global solar
$7.1^{+0.6}_{-0.5}$	<sup>35</sup> SMY	04	FIT KamLAND + global solar
$6.0^{+1.7}_{-1.6}$	<sup>36</sup> SMY	04	FIT global solar
$6.0^{+2.5}_{-1.6}$	<sup>37</sup> SMY	04	FIT SKAM + SNO
2.8–12.0	<sup>38</sup> AHMAD	02b	FIT global solar
3.2–19.1	<sup>39</sup> FUKUDA	02	FIT global solar

<sup>1</sup> GANDO 13 obtained this result by a three-neutrino oscillation analysis using KamLAND, global solar neutrino, short-baseline (SBL) reactor, and accelerator data, assuming *CPT* invariance. Supersedes GANDO 11.

<sup>2</sup> FORERO 14 performs a global fit to  $\Delta m_{21}^2$  using solar, reactor, long-baseline accelerator, and atmospheric neutrino data.

<sup>3</sup> GONZALEZ-GARCIA 14 result comes from a frequentist global fit. The corresponding Bayesian global fit to the same data results are reported in BERGSTROM 15 as  $(7.50^{+0.19}_{-0.17}) \times 10^{-5} \text{ eV}^2$  for normal and  $(7.50^{+0.18}_{-0.17}) \times 10^{-5} \text{ eV}^2$  for inverted mass ordering.

<sup>4</sup> AHARMIM 13 obtained this result by a two-neutrino oscillation analysis using global solar neutrino data.

<sup>5</sup> AHARMIM 13 global solar neutrino data include SNO's all-phases-combined analysis results on the total active  $^8\text{B}$  neutrino flux and energy-dependent  $\nu_e$  survival probability parameters, measurements of Cl (CLEVELAND 98), Ga (ABDURASHITOV 09) which contains combined analysis with GNO (ALTMANN 05 and Ph.D. thesis of F. Kaether), and  $^7\text{Be}$  (BELLINI 11A) rates, and  $^8\text{B}$  solar-neutrino recoil electron measurements of SK-I (HOSAKA 06) zenith, SK-II (CRAVENS 08), and SK-III (ABE 11) day/night spectra, and Borexino (BELLINI 10A) spectra.

<sup>6</sup> AHARMIM 13 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.45 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data.

<sup>7</sup> AHARMIM 13 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.45 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino and KamLAND data (GANDO 11). *CPT* invariance is assumed.

<sup>8</sup> GANDO 13 obtained this result by a three-neutrino oscillation analysis using KamLAND and global solar neutrino data, assuming *CPT* invariance. Supersedes GANDO 11.

<sup>9</sup> GANDO 13 obtained this result by a three-neutrino oscillation analysis using KamLAND data. Supersedes GANDO 11.

<sup>10</sup> ABE 11 obtained this result by a two-neutrino oscillation analysis using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, SAGE, and KamLAND data. *CPT* invariance is assumed.

<sup>11</sup> ABE 11 obtained this result by a two-neutrino oscillation analysis using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, and SAGE data.

<sup>12</sup> ABE 11 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.4 \times 10^{-3} \text{ eV}^2$ , using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, SAGE, and KamLAND data. The normal neutrino mass ordering and *CPT* invariance are assumed.

<sup>13</sup> ABE 11 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.4 \times 10^{-3} \text{ eV}^2$ , using solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, and GALLEX/GNO data. The normal neutrino mass ordering is assumed.

# Lepton Particle Listings

## Neutrino Mixing

- <sup>14</sup> BELLINI 11A obtained this result by a two-neutrino oscillation analysis using KamLAND, Homestake, SAGE, Gallex, GNO, Kamiokande, Super-Kamiokande, SNO, and Borexino (BELLINI 11A) data and the SSM flux prediction in SERENELLI 11 (Astrophysical Journal **743** 24 (2011)) with the exception that the  $^8\text{B}$  flux was left free. CPT invariance is assumed.
- <sup>15</sup> BELLINI 11A obtained this result by a two-neutrino oscillation analysis using Homestake, SAGE, Gallex, GNO, Kamiokande, Super-Kamiokande, SNO, and Borexino (BELLINI 11A) data and the SSM flux prediction in SERENELLI 11 (Astrophysical Journal **743** 24 (2011)) with the exception that the  $^8\text{B}$  flux was left free.
- <sup>16</sup> GANDO 11 obtain this result with three-neutrino fit using the KamLAND + solar data. Superseded by GANDO 13.
- <sup>17</sup> GANDO 11 obtain this result with three-neutrino fit using the KamLAND data only. Supersedes ABE 08A.
- <sup>18</sup> AHARMIM 10 global solar neutrino data include SNO's low-energy-threshold analysis survival probability day/night curves, SNO Phase III integral rates (AHARMIM 08), CL (CLEVELAND 98), SAGE (ABDURASHITOV 09), Gallex/GNO (HAMPEL 99, ALTMANN 05), Borexino (ARPESELLA 08A), SK-I zenith (HOSAKA 06), and SK-II day/night spectra (CRAVENS 08).
- <sup>19</sup> AHARMIM 10 obtained this result by a two-neutrino oscillation analysis using global solar neutrino data and KamLAND data (ABE 08A). CPT invariance is assumed.
- <sup>20</sup> AHARMIM 10 obtained this result by a two-neutrino oscillation analysis using global solar neutrino data.
- <sup>21</sup> AHARMIM 10 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.3 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data and KamLAND data (ABE 08A). CPT invariance is assumed.
- <sup>22</sup> AHARMIM 10 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.3 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data.
- <sup>23</sup> ABE 08A obtained this result by a rate + shape + time combined geoneutrino and reactor two-neutrino fit for  $\Delta m_{21}^2$  and  $\tan^2 \theta_{12}$ , using KamLAND data only. Superseded by GANDO 11.
- <sup>24</sup> ABE 08A obtained this result by means of a two-neutrino fit using KamLAND, Homestake, SAGE, GALLEX, GNO, SK (zenith angle and E-spectrum), the SNO  $\chi^2$ -map, and solar flux data. CPT invariance is assumed. Superseded by GANDO 11.
- <sup>25</sup> AHARMIM 08 obtained this result by a two-neutrino oscillation analysis using all solar neutrino data including those of Borexino (ARPESELLA 08A) and Super-Kamiokande-I (HOSAKA 06), and KamLAND data (ABE 08A). CPT invariance is assumed.
- <sup>26</sup> HOSAKA 06 obtained this result by a two-neutrino oscillation analysis using solar neutrino and KamLAND data (ARAKI 05). CPT invariance is assumed.
- <sup>27</sup> HOSAKA 06 obtained this result by a two-neutrino oscillation analysis using the data from Super-Kamiokande, SNO (AHMAD 02 and AHMAD 02b), and KamLAND (ARAKI 05) experiments. CPT invariance is assumed.
- <sup>28</sup> HOSAKA 06 obtained this result by a two-neutrino oscillation analysis using the Super-Kamiokande and SNO (AHMAD 02 and AHMAD 02b) solar neutrino data.
- <sup>29</sup> HOSAKA 06 obtained this result from the consistency between the observed and expected day-night flux asymmetry amplitude. The listed 68% CL range is derived from the  $1\sigma$  boundary of the amplitude fit to the data. Oscillation parameters are constrained to be in the LMA region. The mixing angle is fixed at  $\tan^2 \theta = 0.44$  because the fit depends only very weakly on it.
- <sup>30</sup> AHARMIM 05A obtained this result by a two-neutrino oscillation analysis using solar neutrino and KamLAND data (ARAKI 05). CPT invariance is assumed. AHARMIM 05A also quotes  $\Delta(m^2) = (8.0^{+0.6}_{-0.4}) \times 10^{-5} \text{ eV}^2$  as the error enveloping the 68% CL two-dimensional region.
- <sup>31</sup> AHARMIM 05A obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in figure 35a of AHARMIM 05A. AHARMIM 05A also quotes  $\Delta(m^2) = (6.5^{+1.4}_{-2.3}) \times 10^{-5} \text{ eV}^2$  as the error enveloping the 68% CL two-dimensional region.
- <sup>32</sup> ARAKI 05 obtained this result by a two-neutrino oscillation analysis using KamLAND and solar neutrino data. CPT invariance is assumed. The  $1\sigma$  error shown here is provided by the KamLAND collaboration. The error quoted in ARAKI 05,  $\Delta(m^2) = (7.9^{+0.6}_{-0.5}) \times 10^{-5}$ , envelops the 68% CL two-dimensional region.
- <sup>33</sup> AHMED 04A obtained this result by a two-neutrino oscillation analysis using solar neutrino and KamLAND data (EGUCHI 03). CPT invariance is assumed. AHMED 04A also quotes  $\Delta(m^2) = (7.1^{+1.2}_{-0.6}) \times 10^{-5} \text{ eV}^2$  as the error enveloping the 68% CL two-dimensional region.
- <sup>34</sup> AHMED 04A obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in Fig. 4(b) of AHMED 04A. The best-fit point is  $\Delta(m^2) = 5.0 \times 10^{-5} \text{ eV}^2$  and  $\tan \theta = 0.34$  ( $\sin^2 2\theta = 0.76$ ).
- <sup>35</sup> SMY 04 obtained this result by a two-neutrino oscillation analysis using solar neutrino and KamLAND data (IANNI 03). CPT invariance is assumed.
- <sup>36</sup> SMY 04 obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The  $1\sigma$  errors are read from Fig. 6(a) of SMY 04.
- <sup>37</sup> SMY 04 obtained this result by a two-neutrino oscillation analysis using the Super-Kamiokande and SNO (AHMAD 02 and AHMAD 02b) solar neutrino data. The  $1\sigma$  errors are read from Fig. 6(a) of SMY 04.
- <sup>38</sup> AHMAD 02b obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95% CL two-dimensional region shown in Fig. 4(b) of AHMAD 02b. The best fit point is  $\Delta(m^2) = 5.0 \times 10^{-5} \text{ eV}^2$  and  $\tan \theta = 0.34$  ( $\sin^2 2\theta = 0.76$ ).
- <sup>39</sup> FUKUDA 02 obtained this result by a two-neutrino oscillation analysis using the data from all solar neutrino experiments. The listed range of the parameter envelops the 95%

CL two-dimensional region shown in Fig. 4 of FUKUDA 02. The best fit point is  $\Delta(m^2) = 6.9 \times 10^{-5} \text{ eV}^2$  and  $\tan^2 \theta = 0.38$  ( $\sin^2 2\theta = 0.80$ ).

### $\sin^2(\theta_{23})$

The reported limits below correspond to the projection onto the  $\sin^2(\theta_{23})$  axis of the 90% CL contours in the  $\sin^2(\theta_{23}) - \Delta m_{32}^2$  plane presented by the authors. Unless otherwise specified, the limits are 90% CL and the reported uncertainties are 68% CL.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.50 ± 0.05</b>	<b>OUR FIT</b>		Assuming inverted mass hierarchy
<b>0.51 ± 0.05</b>	<b>OUR FIT</b>		Assuming normal mass hierarchy
0.53 $^{+0.09}_{-0.12}$	1 AARTSEN	15A ICCB	$3\nu$ osc; normal mass ordering
0.51 $^{+0.09}_{-0.11}$	1 AARTSEN	15A ICCB	$3\nu$ osc; inverted mass ordering
0.514 $^{+0.055}_{-0.056}$	2 ABE	14 T2K	$3\nu$ osc.; normal mass ordering
0.511 ± 0.055	2 ABE	14 T2K	$3\nu$ osc.; inverted mass ordering
0.41 $^{+0.23}_{-0.06}$	3 ADAMSON	14 MINS	$3\nu$ osc., normal mass ordering
0.41 $^{+0.26}_{-0.07}$	3 ADAMSON	14 MINS	$3\nu$ osc.; inverted mass ordering
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.567 $^{+0.032}_{-0.128}$	4 FORERO	14 FIT	Normal mass ordering
0.573 $^{+0.025}_{-0.043}$	4 FORERO	14 FIT	Inverted mass ordering
0.452 $^{+0.052}_{-0.028}$	5 GONZALEZ-G.	14 FIT	Normal mass ordering; global fit
0.579 $^{+0.025}_{-0.037}$	5 GONZALEZ-G.	14 FIT	Inverted mass ordering; global fit
0.24 to 0.76	6 AARTSEN	13B ICCB	DeepCore, $2\nu$ oscillation
0.514 ± 0.082	7 ABE	13G T2K	$3\nu$ osc.; normal mass ordering
0.388 $^{+0.051}_{-0.053}$	8 ADAMSON	13B MINS	Beam + Atmospheric; identical $\nu$ & $\bar{\nu}$
0.3 to 0.7	9 ABE	12A T2K	off-axis beam
0.28 to 0.72	10 ADAMSON	12 MINS	$\bar{\nu}$ beam
0.25 to 0.75	11,12 ADAMSON	12B MINS	MINOS atmospheric
0.27 to 0.73	11,13 ADAMSON	12B MINS	MINOS pure atmospheric $\nu$
0.21 to 0.79	11,13 ADAMSON	12B MINS	MINOS pure atmospheric $\bar{\nu}$
0.15 to 0.85	14 ADRIAN-MAR.	12 ANTR	atmospheric $\nu$ with deep sea telescope
0.39 to 0.61	15 ABE	11C SKAM	Super-Kamiokande
0.34 to 0.66	ADAMSON	11 MINS	$2\nu$ osc.; maximal mixing
0.31 $^{+0.10}_{-0.07}$	16 ADAMSON	11B MINS	$\bar{\nu}$ beam
0.41 to 0.59	17 WENDELL	10 SKAM	$3\nu$ osc. with solar terms; $\theta_{13}=0$
0.39 to 0.61	18 WENDELL	10 SKAM	$3\nu$ osc.; normal mass ordering
0.37 to 0.63	19 WENDELL	10 SKAM	$3\nu$ osc.; inverted mass ordering
0.31 to 0.69	ADAMSON	08A MINS	MINOS
0.05 to 0.95	20 ADAMSON	06 MINS	atmospheric $\nu$ with far detector
0.18 to 0.82	21 AHN	06A K2K	KEK to Super-K
0.23 to 0.77	22 MICHAEL	06 MINS	MINOS
0.18 to 0.82	23 ALIU	05 K2K	KEK to Super-K
0.18 to 0.82	24 ALLISON	05 SOU2	
0.36 to 0.64	25 ASHIE	05 SKAM	Super-Kamiokande
0.28 to 0.72	26 AMBROSIO	04 MCRO	MACRO
0.34 to 0.66	27 ASHIE	04 SKAM	L/E distribution
0.08 to 0.92	28 AHN	03 K2K	KEK to Super-K
0.13 to 0.87	29 AMBROSIO	03 MCRO	MACRO
0.26 to 0.74	30 AMBROSIO	03 MCRO	MACRO
0.15 to 0.85	31 SANCHEZ	03 SOU2	Soudan-2 Atmospheric
0.28 to 0.72	32 AMBROSIO	01 MCRO	upward $\mu$
0.29 to 0.71	33 AMBROSIO	01 MCRO	upward $\mu$
0.13 to 0.87	34 FUKUDA	99c SKAM	upward $\mu$
0.23 to 0.77	35 FUKUDA	99d SKAM	upward $\mu$
0.08 to 0.92	36 FUKUDA	99d SKAM	stop $\mu$ / through
0.29 to 0.71	37 FUKUDA	98c SKAM	Super-Kamiokande
0.08 to 0.92	38 HATA KEYAMA	98 KAMI	Kamiokande
0.24 to 0.76	39 HATA KEYAMA	98 KAMI	Kamiokande
0.20 to 0.80	40 FUKUDA	94 KAMI	Kamiokande

<sup>1</sup> AARTSEN 15A obtains this result by a three-neutrino oscillation analysis using 10–100 GeV muon neutrino sample from a total of 953 days of measurement with the low-energy subdetector DeepCore of the IceCube neutrino telescope.

<sup>2</sup> ABE 14 results are based on  $\nu_\mu$  disappearance using three-neutrino oscillation fit. The confidence intervals are derived from one dimensional profiled likelihoods.

<sup>3</sup> ADAMSON 14 uses a complete set of accelerator and atmospheric data. The analysis combines the  $\nu_\mu$  disappearance and  $\nu_e$  appearance data using three-neutrino oscillation fit. The fit results are obtained for normal and inverted mass ordering assumptions. The best fit is for lower  $\theta_{23}$  quadrant and inverted mass ordering.

<sup>4</sup> FORERO 14 performs a global fit to neutrino oscillations using solar, reactor, long-baseline accelerator, and atmospheric neutrino data.

<sup>5</sup> GONZALEZ-GARCIA 14 result comes from a frequentist global fit. The corresponding Bayesian global fit to the same data results are reported in BERGSTROM 15 as 68% CL intervals of 0.433–0.496 or 0.530–0.594 for normal and 0.514–0.612 for inverted mass ordering.

<sup>6</sup> AARTSEN 13B obtained this result by a two-neutrino oscillation analysis using 20–100 GeV muon neutrino sample from a total of 318.9 days of live-time measurement with the low-energy subdetector DeepCore of the IceCube neutrino telescope.

<sup>7</sup> The best fit value is  $\sin^2(\theta_{23}) = 0.514 \pm 0.082$ . Superseded by ABE 14.

<sup>8</sup> ADAMSON 13b obtained this result from  $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance using  $\nu_\mu$  ( $10.71 \times 10^{20}$  POT) and  $\bar{\nu}_\mu$  ( $3.36 \times 10^{20}$  POT) beams, and atmospheric (37.88kton-years) data from MINOS. The fit assumed two-flavor neutrino hypothesis and identical  $\nu_\mu$  and  $\bar{\nu}_\mu$  oscillation parameters. Superseded by ADAMSON 14.

<sup>9</sup> ABE 12a obtained this result by a two-neutrino oscillation analysis. The best-fit point is  $\sin^2(2\theta_{23}) = 0.98$ .

<sup>10</sup> ADAMSON 12 is a two-neutrino oscillation analysis using antineutrinos. The best fit value is  $\sin^2(2\theta_{23}) = 0.95^{+0.10}_{-0.11} \pm 0.01$ .

<sup>11</sup> ADAMSON 12b obtained this result by a two-neutrino oscillation analysis of the L/E distribution using 37.9 kton-yr atmospheric neutrino data with the MINOS far detector.

<sup>12</sup> The best fit point is  $\Delta m^2 = 0.0019 \text{ eV}^2$  and  $\sin^2 2\theta = 0.99$ . The 90% single-parameter confidence interval at the best fit point is  $\sin^2 2\theta > 0.86$ .

<sup>13</sup> The data are separated into pure samples of  $\nu$ s and  $\bar{\nu}$ s, and separate oscillation parameters for  $\nu$ s and  $\bar{\nu}$ s are fit to the data. The best fit point is  $(\Delta m^2, \sin^2 2\theta) = (0.0022 \text{ eV}^2, 0.99)$  and  $(\Delta \bar{m}^2, \sin^2 2\bar{\theta}) = (0.0016 \text{ eV}^2, 1.00)$ . The quoted result is taken from the 90% C.L. contour in the  $(\Delta m^2, \sin^2 2\theta)$  plane obtained by minimizing the four parameter log-likelihood function with respect to the other oscillation parameters.

<sup>14</sup> ADRIAN-MARTINEZ 12 measured the oscillation parameters of atmospheric neutrinos with the ANTARES deep sea neutrino telescope using the data taken from 2007 to 2010 (863 days of total live time).

<sup>15</sup> ABE 11c obtained this result by a two-neutrino oscillation analysis using the Super-Kamiokande-I+II+III atmospheric neutrino data. ABE 11c also reported results under a two-neutrino disappearance model with separate mixing parameters between  $\nu$  and  $\bar{\nu}$ , and obtained  $\sin^2 2\theta > 0.93$  for  $\nu$  and  $\sin^2 2\theta > 0.83$  for  $\bar{\nu}$  at 90% C.L.

<sup>16</sup> ADAMSON 11b obtained this result by a two-neutrino oscillation analysis of antineutrinos in an antineutrino enhanced beam with  $1.71 \times 10^{20}$  protons on target. This result is consistent with the neutrino measurements of ADAMSON 11 at 2% C.L.

<sup>17</sup> WENDELL 10 obtained this result ( $\sin^2 \theta_{23} = 0.407\text{--}0.583$ ) by a three-neutrino oscillation analysis using the Super-Kamiokande-I+II+III atmospheric neutrino data, assuming  $\theta_{13} = 0$  but including the solar oscillation parameters  $\Delta m_{21}^2$  and  $\sin^2 \theta_{12}$  in the fit.

<sup>18</sup> WENDELL 10 obtained this result ( $\sin^2 \theta_{23} = 0.43\text{--}0.61$ ) by a three-neutrino oscillation analysis with one mass scale dominance ( $\Delta m_{21}^2 = 0$ ) using the Super-Kamiokande-I+II+III atmospheric neutrino data, and updates the HOSAKA 06a result.

<sup>19</sup> WENDELL 10 obtained this result ( $\sin^2 \theta_{23} = 0.44\text{--}0.63$ ) by a three-neutrino oscillation analysis with one mass scale dominance ( $\Delta m_{21}^2 = 0$ ) using the Super-Kamiokande-I+II+III atmospheric neutrino data, and updates the HOSAKA 06a result.

<sup>20</sup> ADAMSON 06 obtained this result by a two-neutrino oscillation analysis of the L/E distribution using 4.54 kton yr atmospheric neutrino data with the MINOS far detector.

<sup>21</sup> Supercedes ALIU 05.

<sup>22</sup> MICHAEL 06 best fit is for maximal mixing. See also ADAMSON 08.

<sup>23</sup> The best fit is for maximal mixing.

<sup>24</sup> ALLISON 05 result is based upon atmospheric neutrino interactions including upward-stopping muons, with an exposure of 5.9 kton yr. From a two-flavor oscillation analysis the best-fit point is  $\Delta m^2 = 0.0017 \text{ eV}^2$  and  $\sin^2(2\theta) = 0.97$ .

<sup>25</sup> ASHIE 05 obtained this result by a two-neutrino oscillation analysis using 92 kton yr atmospheric neutrino data from the complete Super-Kamiokande I running period.

<sup>26</sup> AMBROSIO 04 obtained this result, without using the absolute normalization of the neutrino flux, by combining the angular distribution of upward through-going muon tracks with  $E_\mu > 1 \text{ GeV}$ ,  $N_{\text{low}}$  and  $N_{\text{high}}$ , and the numbers of InDown + UpStop and InUp events. Here,  $N_{\text{low}}$  and  $N_{\text{high}}$  are the number of events with reconstructed neutrino energies  $< 30 \text{ GeV}$  and  $> 130 \text{ GeV}$ , respectively. InDown and InUp represent events with downward and upward-going tracks starting inside the detector due to neutrino interactions, while UpStop represents entering upward-going tracks which stop in the detector. The best fit is for maximal mixing.

<sup>27</sup> ASHIE 04 obtained this result from the L(flight length)/E(estimated neutrino energy) distribution of  $\nu_\mu$  disappearance probability, using the Super-Kamiokande-I 1489 live-day atmospheric neutrino data.

<sup>28</sup> There are several islands of allowed region from this K2K analysis, extending to high values of  $\Delta m^2$ . We only include the one that overlaps atmospheric neutrino analyses. The best fit is for maximal mixing.

<sup>29</sup> AMBROSIO 03 obtained this result on the basis of the ratio  $R = N_{\text{low}}/N_{\text{high}}$ , where  $N_{\text{low}}$  and  $N_{\text{high}}$  are the number of upward through-going muon events with reconstructed neutrino energy  $< 30 \text{ GeV}$  and  $> 130 \text{ GeV}$ , respectively. The data came from the full detector run started in 1994. The method of FELDMAN 98 is used to obtain the limits.

<sup>30</sup> AMBROSIO 03 obtained this result by using the ratio R and the angular distribution of the upward through-going muons. R is given in the previous note and the angular distribution is reported in AMBROSIO 01. The method of FELDMAN 98 is used to obtain the limits. The best fit is to maximal mixing.

<sup>31</sup> SANCHEZ 03 is based on an exposure of 5.9 kton yr. The result is obtained using a likelihood analysis of the neutrino L/E distribution for a selection  $\mu$  flavor sample while the e-flavor sample provides flux normalization. The method of FELDMAN 98 is used to obtain the allowed region. The best fit is  $\sin^2(2\theta) = 0.97$ .

<sup>32</sup> AMBROSIO 01 result is based on the angular distribution of upward through-going muon tracks with  $E_\mu > 1 \text{ GeV}$ . The data came from three different detector configurations, but the statistics is largely dominated by the full detector run, from May 1994 to December 2000. The total live time, normalized to the full detector configuration is 6.17 years. The best fit is obtained outside the physical region. The method of FELDMAN 98 is used to obtain the limits. The best fit is for maximal mixing.

<sup>33</sup> AMBROSIO 01 result is based on the angular distribution and normalization of upward through-going muon tracks with  $E_\mu > 1 \text{ GeV}$ . See the previous footnote.

<sup>34</sup> FUKUDA 99c obtained this result from a total of 537 live days of upward through-going muon data in Super-Kamiokande between April 1996 to January 1998. With a threshold of  $E_\mu > 1.6 \text{ GeV}$ , the observed flux is  $(1.74 \pm 0.07 \pm 0.02) \times 10^{-13} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$ . The best fit is  $\sin^2(2\theta) = 0.95$ .

<sup>35</sup> FUKUDA 99d obtained this result from a simultaneous fitting to zenith angle distributions of upward-stopping and through-going muons. The flux of upward-stopping muons of minimum energy of 1.6 GeV measured between April 1996 and January 1998 is  $(0.39 \pm 0.04 \pm 0.02) \times 10^{-13} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$ . This is compared to the expected flux of  $(0.73 \pm 0.16 \text{ (theoretical error)}) \times 10^{-13} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$ . The best fit is to maximal mixing.

<sup>36</sup> FUKUDA 99d obtained this result from the zenith dependence of the upward-stopping/through-going flux ratio. The best fit is to maximal mixing.

<sup>37</sup> FUKUDA 98c obtained this result by an analysis of 33.0 kton yr atmospheric neutrino data. The best fit is for maximal mixing.

<sup>38</sup> HATAKEYAMA 98 obtained this result from a total of 2456 live days of upward-going muon data in Kamiokande between December 1985 and May 1995. With a threshold of  $E_\mu > 1.6 \text{ GeV}$ , the observed flux of upward through-going muons is  $(1.94 \pm 0.10^{+0.07}_{-0.06}) \times 10^{-13} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$ . This is compared to the expected flux of  $(2.46 \pm 0.54 \text{ (theoretical error)}) \times 10^{-13} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$ . The best fit is for maximal mixing.

<sup>39</sup> HATAKEYAMA 98 obtained this result from a combined analysis of Kamiokande contained events (FUKUDA 94) and upward going muon events. The best fit is  $\sin^2(2\theta) = 0.95$ .

<sup>40</sup> FUKUDA 94 obtained the result by a combined analysis of sub- and multi-GeV atmospheric neutrino events in Kamiokande. The best fit is for maximal mixing.

### $\Delta m_{32}^2$

The sign of  $\Delta m_{32}^2$  is not known at this time. Only the absolute value is quoted below. Unless otherwise specified, the ranges below correspond to the projection onto the  $\Delta m_{32}^2$  axis of the 90% CL contours in the  $\sin^2(2\theta_{23}) - \Delta m_{32}^2$  plane presented by the authors. If uncertainties are reported with the value, they correspond to one standard deviation uncertainty.

VALUE ( $10^{-3} \text{ eV}^2$ )	DOCUMENT ID	TECN	COMMENT
<b>2.51 <math>\pm</math> 0.06 OUR FIT</b>			Assuming inverted mass hierarchy
<b>2.44 <math>\pm</math> 0.06 OUR FIT</b>			Assuming normal mass hierarchy
2.72 $^{+0.19}_{-0.20}$	1 AARTSEN	15A ICCB	3 $\nu$ osc; normal mass ordering
2.73 $^{+0.18}_{-0.21}$	1 AARTSEN	15A ICCB	3 $\nu$ osc; inverted mass ordering
2.37 $\pm$ 0.11	2 AN	15 DAYA	3 $\nu$ osc.; normal mass ordering
2.47 $\pm$ 0.11	2 AN	15 DAYA	3 $\nu$ osc.; inverted mass ordering
2.51 $\pm$ 0.10	3 ABE	14 T2K	3 $\nu$ osc.; normal mass ordering
2.56 $\pm$ 0.10	3 ABE	14 T2K	3 $\nu$ osc.; inverted mass ordering
2.37 $\pm$ 0.09	4 ADAMSON	14 MINS	3 $\nu$ osc., accel., atmospheric; normal mass ordering
2.41 $^{+0.12}_{-0.09}$	4 ADAMSON	14 MINS	3 $\nu$ osc., accel., atmospheric; inverted mass ordering
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.54 $^{+0.19}_{-0.20}$	5 AN	14 DAYA	3 $\nu$ osc.; normal mass ordering
2.64 $^{+0.19}_{-0.20}$	5 AN	14 DAYA	3 $\nu$ osc.; inverted mass ordering
2.48 $^{+0.05}_{-0.07}$	6 FORERO	14 FIT	3 $\nu$ ; normal mass ordering
2.38 $^{+0.05}_{-0.06}$	6 FORERO	14 FIT	3 $\nu$ ; inverted mass ordering
2.457 $\pm$ 0.047	7,8 GONZALEZ-G..14	FIT	Normal mass ordering; global fit
2.449 $^{+0.048}_{-0.047}$	7 GONZALEZ-G..14	FIT	Inverted mass ordering; global fit
2.3 $^{+0.6}_{-0.5}$	9 AARTSEN	13B ICCB	DeepCore, 2 $\nu$ oscillation
2.44 $^{+0.17}_{-0.15}$	10 ABE	13G T2K	3 $\nu$ osc.; normal mass ordering
2.41 $^{+0.09}_{-0.10}$	11 ADAMSON	13B MINS	2 $\nu$ osc.; beam + atmospheric; identical $\nu$ & $\bar{\nu}$
2.2–3.1	12 ABE	12A T2K	off-axis beam
2.62 $^{+0.31}_{-0.28} \pm 0.09$	13 ADAMSON	12 MINS	$\bar{\nu}$ beam
1.35–2.55	14,15 ADAMSON	12B MINS	MINOS atmospheric
1.4–5.6	14,16 ADAMSON	12B MINS	MINOS pure atmospheric $\nu$
0.9–2.5	14,16 ADAMSON	12B MINS	MINOS pure atmospheric $\bar{\nu}$
1.8–5.0	17 ADRIAN-MAR.12	ANTR	atm. $\nu$ with deep see telescope
1.3–4.0	18 ABE	11C SKAM	atmospheric $\bar{\nu}$
2.32 $^{+0.12}_{-0.08}$	ADAMSON	11 MINS	2 $\nu$ oscillation; maximal mixing
3.36 $^{+0.46}_{-0.40}$	19 ADAMSON	11B MINS	$\bar{\nu}$ beam
<3.37	20 ADAMSON	11C MINS	MINOS
1.9–2.6	21 WENDELL	10 SKAM	3 $\nu$ osc.; normal mass ordering
1.7–2.7	21 WENDELL	10 SKAM	3 $\nu$ osc.; inverted mass ordering
2.43 $\pm$ 0.13	ADAMSON	08A MINS	MINOS
0.07–50	22 ADAMSON	06 MINS	atmospheric $\nu$ with far detector
1.9–4.0	23,24 AHN	06A K2K	KEK to Super-K
2.2–3.8	25 MICHAEL	06 MINS	MINOS

# Lepton Particle Listings

## Neutrino Mixing

- 1.9–3.6  
0.3–12  
1.5–3.4  
0.6–8.0  
1.9 to 3.0  
1.5–3.9  
0.25–9.0  
0.6–7.0  
0.15–15  
0.6–15  
1.0–6.0  
1.0–50  
1.5–15.0  
0.7–18  
0.5–6.0  
0.55–50  
4–23  
5–25
- 23 ALIU 05 K2K KEK to Super-K  
26 ALLISON 05 SOU2  
27 ASHIE 05 SKAM atmospheric neutrino  
28 AMBROSIO 04 MCRO MACRO  
29 ASHIE 04 SKAM L/E distribution  
30 AHN 03 K2K KEK to Super-K  
31 AMBROSIO 03 MCRO MACRO  
32 AMBROSIO 03 MCRO MACRO  
33 SANCHEZ 03 SOU2 Soudan-2 Atmospheric  
34 AMBROSIO 01 MCRO upward  $\mu$   
35 AMBROSIO 01 MCRO upward  $\mu$   
36 FUKUDA 99c SKAM upward  $\mu$   
37 FUKUDA 99d SKAM upward  $\mu$   
38 FUKUDA 99d SKAM stop  $\mu$  / through  
39 FUKUDA 98c SKAM Super-Kamiokande  
40 HATAKEYAMA 98 KAMI Kamiokande  
41 HATAKEYAMA 98 KAMI Kamiokande  
42 FUKUDA 94 KAMI Kamiokande
- 1 AARTSEN 15A obtains this result by a three-neutrino oscillation analysis using 10–100 GeV muon neutrino sample from a total of 953 days of measurements with the low-energy subdetector DeepCore of the IceCube neutrino telescope.
- 2 AN 15 uses all eight identical detectors, with four placed near the reactor cores and the remaining four at the far hall to determine prompt energy spectra. The results correspond to the exposure of  $6.9 \times 10^5$  GW<sub>th</sub>-ton-days. They derive  $\Delta m^2 = (2.42 \pm 0.11) \times 10^{-3}$  eV<sup>2</sup>. Assuming the normal (inverted) ordering, the fitted  $\Delta m^2_{32} = (2.37 \pm 0.11) \times 10^{-3}$  (( $2.47 \pm 0.11$ )  $\times 10^{-3}$ ) eV<sup>2</sup>. Supersedes AN 14.
- 3 ABE 14 results are based on  $\nu_\mu$  disappearance using three-neutrino oscillation fit. The confidence intervals are derived from one dimensional profiled likelihoods. In ABE 14 the inverted mass ordering result is reported as  $\Delta m^2_{13} = (2.48 \pm 0.10) \times 10^{-3}$  eV<sup>2</sup> which we converted to  $\Delta m^2_{32}$  by adding PDG 14 value of  $\Delta m^2_{21} = (7.53 \pm 0.18) \times 10^{-5}$  eV<sup>2</sup>.
- 4 ADAMSON 14 uses a complete set of accelerator and atmospheric data. The analysis combines the analysis combines the  $\nu_\mu$  disappearance and  $\nu_e$  appearance data using three-neutrino oscillation fit. The fit results are obtained for normal and inverted mass ordering assumptions.
- 5 AN 14 uses six identical detectors, with three placed near the reactor cores (flux-weighted baselines of 512 and 561 m) and the remaining three at the far hall (at the flux averaged distance of 1579 m from all six reactor cores) to determine prompt energy spectra and derive  $\Delta m^2_{ee} = (2.59 \pm 0.19) \times 10^{-3}$  eV<sup>2</sup>. Assuming the normal (inverted) ordering, the fitted  $\Delta m^2_{32} = (2.54 \pm 0.19) \times 10^{-3}$  (( $2.64 \pm 0.19$ )  $\times 10^{-3}$ ) eV<sup>2</sup>. Superseded by AN 15.
- 6 FORERO 14 performs a global fit to  $\Delta m^2_{31}$  using solar, reactor, long-baseline accelerator, and atmospheric neutrino data.
- 7 GONZALEZ-GARCIA 14 result comes from a frequentist global fit. The corresponding Bayesian global fit to the same data results are reported in BERGSTROM 15 as ( $2.460 \pm 0.046$ )  $\times 10^{-3}$  eV<sup>2</sup> for normal and ( $2.445 \pm 0.047$ )  $\times 10^{-3}$  eV<sup>2</sup> for inverted mass ordering.
- 8 The value for normal mass ordering is actually a measurement of  $\Delta m^2_{31}$  which differs from  $\Delta m^2_{32}$  by a much smaller value of  $\Delta m^2_{12}$ .
- 9 AARTSEN 13B obtained this result by a two-neutrino oscillation analysis using 20–100 GeV muon neutrino sample from a total of 318.9 days of live-time measurement with the low-energy subdetector DeepCore of the IceCube neutrino telescope.
- 10 Based on the observation of 58  $\nu_\mu$  events with  $205 \pm 17$ (syst) expected in the absence of neutrino oscillations. Superseded by ABE 14.
- 11 ADAMSON 13b obtained this result from  $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance using  $\nu_\mu$  ( $10.71 \times 10^{20}$  POT) and  $\bar{\nu}_\mu$  ( $3.36 \times 10^{20}$  POT) beams, and atmospheric (37.88 kton-years) data from MINOS. The fit assumed two-flavor neutrino hypothesis and identical  $\nu_\mu$  and  $\bar{\nu}_\mu$  oscillation parameters.
- 12 ABE 12A obtained this result by a two-neutrino oscillation analysis. The best-fit point is  $\Delta m^2_{32} = 2.65 \times 10^{-3}$  eV<sup>2</sup>.
- 13 ADAMSON 12 is a two-neutrino oscillation analysis using antineutrinos.
- 14 ADAMSON 12B obtained this result by a two-neutrino oscillation analysis of the L/E distribution using 37.9 kton-yr atmospheric neutrino data with the MINOS far detector.
- 15 The 90% single-parameter confidence interval at the best fit point is  $\Delta m^2 = 0.0019 \pm 0.0004$  eV<sup>2</sup>.
- 16 The data are separated into pure samples of  $\nu_s$  and  $\bar{\nu}_s$ , and separate oscillation parameters for  $\nu_s$  and  $\bar{\nu}_s$  are fit to the data. The best fit point is ( $\Delta m^2, \sin^2 2\theta$ ) = (0.0022 eV<sup>2</sup>, 0.99) and ( $\Delta \bar{m}^2, \sin^2 2\bar{\theta}$ ) = (0.0016 eV<sup>2</sup>, 1.00). The quoted result is taken from the 90% C.L. contour in the ( $\Delta m^2, \sin^2 2\theta$ ) plane obtained by minimizing the four parameter log-likelihood function with respect to the other oscillation parameters.
- 17 ADRIAN-MARTINEZ 12 measured the oscillation parameters of atmospheric neutrinos with the ANTARES deep sea neutrino telescope using the data taken from 2007 to 2010 (863 days of total live time).
- 18 ABE 11c obtained this result by a two-neutrino oscillation analysis with separate mixing parameters between neutrinos and antineutrinos, using the Super-Kamiokande-I+II+III atmospheric neutrino data. The corresponding 90% CL neutrino oscillation parameter range obtained from this analysis is  $\Delta m^2 = 1.7\text{--}3.0 \times 10^{-3}$  eV<sup>2</sup>.
- 19 ADAMSON 11b obtained this result by a two-neutrino oscillation analysis of antineutrinos in an antineutrino enhanced beam with  $1.71 \times 10^{20}$  protons on target. This results is consistent with the neutrino measurements of ADAMSON 11 at 2% C.L.
- 20 ADAMSON 11c obtains this result based on a study of antineutrinos in a neutrino beam and assumes maximal mixing in the two-flavor approximation.
- 21 WENDELL 10 obtained this result by a three-neutrino oscillation analysis with one mass scale dominance ( $\Delta m^2_{21} = 0$ ) using the Super-Kamiokande-I+II+III atmospheric neutrino data, and updates the HOSAKA 06A result.

- 22 ADAMSON 06 obtained this result by a two-neutrino oscillation analysis of the L/E distribution using 4.54 kton yr atmospheric neutrino data with the MINOS far detector.
- 23 The best fit in the physical region is for  $\Delta m^2 = 2.8 \times 10^{-3}$  eV<sup>2</sup>.
- 24 Supersedes ALIU 05.
- 25 MICHAEL 06 best fit is  $2.74 \times 10^{-3}$  eV<sup>2</sup>. See also ADAMSON 08.
- 26 ALLISON 05 result is based on an atmospheric neutrino observation with an exposure of 5.9 kton yr. From a two-flavor oscillation analysis the best-fit point is  $\Delta m^2 = 0.0017$  eV<sup>2</sup> and  $\sin^2 2\theta = 0.97$ .
- 27 ASHIE 05 obtained this result by a two-neutrino oscillation analysis using 92 kton yr atmospheric neutrino data from the complete Super-Kamiokande-I running period. The best fit is for  $\Delta m^2 = 2.1 \times 10^{-3}$  eV<sup>2</sup>.
- 28 AMBROSIO 04 obtained this result, without using the absolute normalization of the neutrino flux, by combining the angular distribution of upward through-going muon tracks with  $E_\mu > 1$  GeV,  $N_{low}$  and  $N_{high}$ , and the numbers of InDown + UpStop and InUp events. Here,  $N_{low}$  and  $N_{high}$  are the number of events with reconstructed neutrino energies  $< 30$  GeV and  $> 130$  GeV, respectively. InDown and InUp represent events with downward and upward-going tracks starting inside the detector due to neutrino interactions, while UpStop represents entering upward-going tracks which stop in the detector. The best fit is for  $\Delta m^2 = 2.3 \times 10^{-3}$  eV<sup>2</sup>.
- 29 ASHIE 04 obtained this result from the L(flight length)/E(estimated neutrino energy) distribution of  $\nu_\mu$  disappearance probability, using the Super-Kamiokande-I 1489 live-day atmospheric neutrino data. The best fit is for  $\Delta m^2 = 2.1 \times 10^{-3}$  eV<sup>2</sup>.
- 30 There are several islands of allowed region from this K2K analysis, extending to high values of  $\Delta m^2$ . We only include the one that overlaps atmospheric neutrino analyses. The best fit is for  $\Delta m^2 = 2.8 \times 10^{-3}$  eV<sup>2</sup>.
- 31 AMBROSIO 03 obtained this result on the basis of the ratio  $R = N_{low}/N_{high}$ , where  $N_{low}$  and  $N_{high}$  are the number of upward through-going muon events with reconstructed neutrino energy  $< 30$  GeV and  $> 130$  GeV, respectively. The data came from the full detector run started in 1994. The method of FELDMAN 98 is used to obtain the limits. The best fit is for  $\Delta m^2 = 2.5 \times 10^{-3}$  eV<sup>2</sup>.
- 32 AMBROSIO 03 obtained this result by using the ratio  $R$  and the angular distribution of the upward through-going muons.  $R$  is given in the previous note and the angular distribution is reported in AMBROSIO 01. The method of FELDMAN 98 is used to obtain the limits. The best fit is for  $\Delta m^2 = 2.5 \times 10^{-3}$  eV<sup>2</sup>.
- 33 SANCHEZ 03 is based on an exposure of 5.9 kton yr. The result is obtained using a likelihood analysis of the neutrino L/E distribution for a selection  $\mu$  flavor sample while the  $e$ -flavor sample provides flux normalization. The method of FELDMAN 98 is used to obtain the allowed region. The best fit is for  $\Delta m^2 = 5.2 \times 10^{-3}$  eV<sup>2</sup>.
- 34 AMBROSIO 01 result is based on the angular distribution of upward through-going muon tracks with  $E_\mu > 1$  GeV. The data came from three different detector configurations, but the statistics is largely dominated by the full detector run, from May 1994 to December 2000. The total live time, normalized to the full detector configuration is 6.17 years. The best fit is obtained outside the physical region. The method of FELDMAN 98 is used to obtain the limits.
- 35 AMBROSIO 01 result is based on the angular distribution and normalization of upward through-going muon tracks with  $E_\mu > 1$  GeV. See the previous footnote.
- 36 FUKUDA 99c obtained this result from a total of 537 live days of upward through-going muon data in Super-Kamiokande between April 1996 to January 1998. With a threshold of  $E_\mu > 1.6$  GeV, the observed flux is  $(1.74 \pm 0.07 \pm 0.02) \times 10^{-13}$  cm<sup>-2</sup>s<sup>-1</sup>sr<sup>-1</sup>. The best fit is for  $\Delta m^2 = 5.9 \times 10^{-3}$  eV<sup>2</sup>.
- 37 FUKUDA 99d obtained this result from a simultaneous fitting to zenith angle distributions of upward-stopping and through-going muons. The flux of upward-stopping muons of minimum energy of 1.6 GeV measured between April 1996 and January 1998 is  $(0.39 \pm 0.04 \pm 0.02) \times 10^{-13}$  cm<sup>-2</sup>s<sup>-1</sup>sr<sup>-1</sup>. This is compared to the expected flux of  $(0.73 \pm 0.16$  (theoretical error))  $\times 10^{-13}$  cm<sup>-2</sup>s<sup>-1</sup>sr<sup>-1</sup>. The best fit is for  $\Delta m^2 = 3.9 \times 10^{-3}$  eV<sup>2</sup>.
- 38 FUKUDA 99d obtained this result from the zenith dependence of the upward-stopping/through-going flux ratio. The best fit is for  $\Delta m^2 = 3.1 \times 10^{-3}$  eV<sup>2</sup>.
- 39 FUKUDA 98c obtained this result by an analysis of 33.0 kton yr atmospheric neutrino data. The best fit is for  $\Delta m^2 = 2.2 \times 10^{-3}$  eV<sup>2</sup>.
- 40 HATAKEYAMA 98 obtained this result from a total of 2456 live days of upward-going muon data in Kamiokande between December 1985 and May 1995. With a threshold of  $E_\mu > 1.6$  GeV, the observed flux of upward through-going muons is  $(1.94 \pm 0.10 \pm 0.07) \times 10^{-13}$  cm<sup>-2</sup>s<sup>-1</sup>sr<sup>-1</sup>. This is compared to the expected flux of  $(2.46 \pm 0.54$  (theoretical error))  $\times 10^{-13}$  cm<sup>-2</sup>s<sup>-1</sup>sr<sup>-1</sup>. The best fit is for  $\Delta m^2 = 2.2 \times 10^{-3}$  eV<sup>2</sup>.
- 41 HATAKEYAMA 98 obtained this result from a combined analysis of Kamiokande contained events (FUKUDA 94) and upward going muon events. The best fit is for  $\Delta m^2 = 13 \times 10^{-3}$  eV<sup>2</sup>.
- 42 FUKUDA 94 obtained the result by a combined analysis of sub- and multi-GeV atmospheric neutrino events in Kamiokande. The best fit is for  $\Delta m^2 = 16 \times 10^{-3}$  eV<sup>2</sup>.

### $\sin^2(\theta_{13})$

At present time direct measurements of  $\sin^2(\theta_{13})$  are derived from the reactor  $\bar{\nu}_e$  disappearance at distances corresponding to the  $\Delta m^2_{32}$  value, i.e.  $L \sim 1$  km. Alternatively, limits can also be obtained from the analysis of the solar neutrino data and accelerator-based  $\nu_\mu \rightarrow \nu_e$  experiments.

VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.19 ± 0.12 OUR AVERAGE</b>				
2.15 ± 0.13	1	AN	15	DAYA DayaBay, Ling Ao/Ao II reactors
2.3 ± 0.9 – 0.8	2	ABE	14H	DCHZ Chooz reactors

See key on page 601

# Lepton Particle Listings

## Neutrino Mixing

$2.12 \pm 0.47$	<sup>3</sup> AN	14B	DAYA	DayaBay, Ling Ao/Ao II reactors
$2.5 \pm 0.9 \pm 0.9$	<sup>4</sup> ABE	13C	DCHZ	Chooz reactors
$2.9 \pm 0.3 \pm 0.5$	<sup>5</sup> AHN	12	RENO	Yonggwang reactors
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.6 \pm 1.2$ $-1.1$	<sup>6</sup> ABE	14A	DCHZ	Chooz reactors
$3.0 \pm 1.3$ $-1.0$	<sup>7</sup> ABE	14C	T2K	Inverted mass ordering
$3.6 \pm 1.0$ $-0.9$	<sup>7</sup> ABE	14C	T2K	Normal mass ordering
$2.3 \pm 0.2$	<sup>8</sup> AN	14	DAYA	DayaBay, Ling Ao/Ao II reactors
$2.34 \pm 0.20$	<sup>9</sup> FORERO	14	FIT	Normal mass ordering
$2.40 \pm 0.19$	<sup>9</sup> FORERO	14	FIT	Inverted mass ordering
$2.18 \pm 0.10$	<sup>10</sup> GONZALEZ-GARCIA	FIT		Normal mass ordering; global fit
$2.19 \pm 0.11$ $-0.10$	<sup>10</sup> GONZALEZ-GARCIA	FIT		Inverted mass ordering; global fit
$2.3 \pm 1.3$ $-1.0$	<sup>11</sup> ABE	13E	T2K	Normal mass ordering
$2.8 \pm 1.6$ $-1.2$	<sup>11</sup> ABE	13E	T2K	Inverted mass ordering
$1.6 \pm 1.3$ $-0.9$	<sup>12</sup> ADAMSON	13A	MINS	Normal mass ordering
$3.0 \pm 1.8$ $-1.6$	<sup>12</sup> ADAMSON	13A	MINS	Inverted mass ordering
<13	<sup>90</sup> AGAFONOVA	13	OPER	OPERA: $3\nu$
<3.6	<sup>95</sup> AHARMIM	13	FIT	global solar: $3\nu$
$2.3 \pm 0.3 \pm 0.1$	<sup>14</sup> AN	13	DAYA	DayaBay, Ling Ao/Ao II reactors
$2.2 \pm 1.1 \pm 0.8$	<sup>15</sup> ABE	12	DCHZ	Chooz reactors
$2.8 \pm 0.8 \pm 0.7$	<sup>16</sup> ABE	12B	DCHZ	Chooz reactors
$2.4 \pm 0.4 \pm 0.1$	<sup>17</sup> AN	12	DAYA	DayaBay, Ling Ao/Ao II reactors
$2.5 \pm 1.8$ $-1.6$	<sup>18</sup> ABE	11	FIT	KamLAND + global solar
<6.1	<sup>95</sup> <sup>19</sup> ABE	11	FIT	Global solar
$1.3$ to $5.6$	<sup>68</sup> <sup>20</sup> ABE	11A	T2K	Normal mass ordering
$1.5$ to $5.6$	<sup>68</sup> <sup>21</sup> ABE	11A	T2K	Inverted mass ordering
$0.3$ to $2.3$	<sup>68</sup> <sup>22</sup> ADAMSON	11D	MINS	Normal mass ordering
$0.8$ to $3.9$	<sup>68</sup> <sup>23</sup> ADAMSON	11D	MINS	Inverted mass ordering
$8 \pm 3$	<sup>68</sup> <sup>24</sup> FOGLI	11	FIT	Global neutrino data
$7.8 \pm 6.2$	<sup>68</sup> <sup>25</sup> GANDO	11	FIT	KamLAND + solar: $3\nu$
$12.4 \pm 13.3$	<sup>68</sup> <sup>26</sup> GANDO	11	FIT	KamLAND: $3\nu$
$3 \pm 9$ $-7$	<sup>90</sup> <sup>27</sup> ADAMSON	10A	MINS	Normal mass ordering
$6 \pm 14$ $-6$	<sup>90</sup> <sup>28</sup> ADAMSON	10A	MINS	Inverted mass ordering
$8 \pm 7$	<sup>29,30</sup> AHARMIM	10	FIT	KamLAND + global solar: $3\nu$
<30	<sup>95</sup> <sup>29,31</sup> AHARMIM	10	FIT	global solar: $3\nu$
<15	<sup>90</sup> <sup>32</sup> WENDELL	10	SKAM	$3\nu$ osc.; normal $m$ ordering
<33	<sup>90</sup> <sup>32</sup> WENDELL	10	SKAM	$3\nu$ osc.; inverted $m$ ordering
$11 \pm 11$ $-8$	<sup>33</sup> ADAMSON	09	MINS	Normal mass ordering
$18 \pm 15$ $-11$	<sup>34</sup> ADAMSON	09	MINS	Inverted mass ordering
$6 \pm 4$	<sup>35</sup> FOGLI	08	FIT	Global neutrino data
$8 \pm 7$	<sup>36</sup> FOGLI	08	FIT	Solar + KamLAND data
$5 \pm 5$	<sup>37</sup> FOGLI	08	FIT	Atmospheric+LBL+CHOOZ
<36	<sup>90</sup> <sup>38</sup> YAMAMOTO	06	K2K	Accelerator experiment
<48	<sup>90</sup> <sup>39</sup> AHN	04	K2K	Accelerator experiment
<36	<sup>90</sup> <sup>40</sup> BOEHM	01		Palo Verde react.
<45	<sup>90</sup> <sup>41</sup> BOEHM	00		Palo Verde react.
<15	<sup>90</sup> <sup>42</sup> APOLLONIO	99	CHOZ	Reactor Experiment

<sup>1</sup> AN 15 uses all eight identical detectors, with four placed near the reactor cores and the remaining four at the far hall to determine the mixing angle  $\theta_{13}$  using the  $\overline{\nu}_e$  observed interaction rates with neutron capture on Gd and energy spectra. The result corresponds to the exposure of  $6.9 \times 10^5$  GW<sub>th</sub>-ton-days. Supersedes AN 14.

<sup>2</sup> ABE 14H uses 467.9 live days of one detector, 1050 m away from two reactor cores of the Chooz nuclear power station, to determine the mixing parameter  $\sin^2(2\theta_{13})$ . The Bugey4 data (DECLAIS 94) is used to constrain the neutrino flux. The data set includes 7.24 reactor-off days. A rate and shape analysis is performed. Supersedes ABE 14A.

<sup>3</sup> AN 14B uses six identical anti-neutrino detectors with flux-weighted baselines of  $\sim 500$  m and  $\sim 1.6$  km to six power reactors. This rate analysis uses a 217-day data set and neutron capture on protons (not Gd) only.  $\Delta m_{31}^2 = 2.32 \times 10^{-3}$  eV<sup>2</sup> is assumed.

<sup>4</sup> ABE 13C uses delayed neutron capture on hydrogen instead of on Gd used previously. The physical volume is thus three times larger. The fit is based on the rate and shape analysis as in ABE 12B. The Bugey4 data (DECLAIS 94) is used to constrain the neutrino flux.

<sup>5</sup> AHN 12 uses two identical detectors, placed at flux weighted distances of 408.56 m and 1433.99 m from six reactor cores, to determine the mixing angle  $\theta_{13}$ . This rate-only analysis excludes the no-oscillation hypothesis at 4.9 standard deviations. The value of  $\Delta m_{31}^2 = (2.32 \pm 0.12 \pm 0.08) \times 10^{-3}$  eV<sup>2</sup> was assumed in the analysis.

<sup>6</sup> ABE 14A uses 467.9 live days of one detector, 1050 m away from two reactor cores of the Chooz nuclear power station, to determine the mixing parameter  $\sin^2(2\theta_{13})$ . The Bugey4 data (DECLAIS 94) is used to constrain the neutrino flux. The data set includes 7.24 reactor-off days. A "rate-modulation" analysis is performed. Supersedes ABE 12B.

<sup>7</sup> ABE 14C result is for  $\nu_e$  appearance and assumes  $\Delta m_{32}^2 = 2.4 \times 10^{-3}$  eV<sup>2</sup>,  $\sin^2(\theta_{23}) = 0.5$ , and  $\delta = 0$ .

<sup>8</sup> AN 14 uses six identical detectors, with three placed near the reactor cores (flux-weighted baselines of 512 and 561 m) and the remaining three at the far hall (at the flux averaged distance of 1579 m from all six reactor cores) to determine the mixing angle  $\theta_{13}$  using the  $\overline{\nu}_e$  observed interaction rates with neutron capture on Gd and energy spectra. Supersedes AN 13 and superseded by AN 15.

<sup>9</sup> FORERO 14 performs a global fit to neutrino oscillations using solar, reactor, long-baseline accelerator, and atmospheric neutrino data.

<sup>10</sup> GONZALEZ-GARCIA 14 result comes from a frequentist global fit. The corresponding Bayesian global fit to the same data results are reported in BERGSTROM 15 as  $(2.18 \pm 0.10 \pm 0.11) \times 10^{-2}$  eV<sup>2</sup> for normal and  $(2.19 \pm 0.12 \pm 0.10) \times 10^{-2}$  eV<sup>2</sup> for inverted mass ordering.

<sup>11</sup> ABE 13E assumes maximal  $\theta_{23}$  mixing and  $CP$  phase  $\delta = 0$ .

<sup>12</sup> ADAMSON 13A results obtained from  $\nu_e$  appearance, assuming  $\delta = 0$ , and  $\sin^2(2\theta_{23}) = 0.957$ .

<sup>13</sup> AHARMIM 13 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.45 \times 10^{-3}$  eV<sup>2</sup>, using global solar neutrino data. AHARMIM 13 global solar neutrino data include SNO's all-phases-combined analysis results on the total active  $^8B$  neutrino flux and energy-dependent  $\nu_e$  survival probability parameters, measurements of Cl (CLEVELAND 98), Ga (ABDURASHITOV 09 which contains combined analysis with GNO (ALTMANN 05 and Ph.D. thesis of F. Kaether)), and  $^7Be$  (BELLINI 11A) rates, and  $^8B$  solar-neutrino recoil electron measurements of SK-I (HOSAKA 06) zenith, SK-II (CRAVENS 08) and SK-III (ABE 11) day/night spectra, and Borexino (BELLINI 10A) spectra. AHARMIM 13 also reported a result combining global solar and KamLAND data, which is  $\sin^2(2\theta_{13}) = (9.1 \pm 2.9 \pm 3.1) \times 10^{-2}$ .

<sup>14</sup> AN 13 uses six identical detectors, with three placed near the reactor cores (flux-weighted baselines of 498 and 555 m) and the remaining three at the far hall (at the flux averaged distance of 1628 m from all six reactor cores) to determine the  $\overline{\nu}_e$  interaction rate ratios. Superseded by AN 14.

<sup>15</sup> ABE 12 determines the  $\overline{\nu}_e$  interaction rate in a single detector, located 1050 m from the cores of two reactors. A rate and shape analysis is performed. The rate normalization is fixed by the results of the Bugey4 reactor experiment, thus avoiding any dependence on possible very short baseline oscillations. The value of  $\Delta m_{31}^2 = 2.4 \times 10^{-3}$  eV<sup>2</sup> is used in the analysis. Superseded by ABE 12B.

<sup>16</sup> ABE 12B determines the neutrino mixing angle  $\theta_{13}$  using a single detector, located 1050 m from the cores of two reactors. This result is based on a spectral shape and rate analysis. The Bugey4 data (DECLAIS 94) is used to constrain the neutrino flux. Superseded by ABE 14A.

<sup>17</sup> AN 12 uses six identical detectors with three placed near the reactor cores (flux-weighted baselines of 470 m and 576 m) and the remaining three at the far hall (at the flux averaged distance of 1648 m from all six reactor cores) to determine the mixing angle  $\theta_{13}$  using the  $\overline{\nu}_e$  observed interaction rate ratios. This rate-only analysis excludes the no-oscillation hypothesis at 5.2 standard deviations. The value of  $\Delta m_{31}^2 = (2.32 \pm 0.12 \pm 0.08) \times 10^{-3}$  eV<sup>2</sup> was assumed in the analysis. Superseded by AN 13.

<sup>18</sup> ABE 11 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.4 \times 10^{-3}$  eV<sup>2</sup>, using global solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, GALLEX/GNO, SAGE, and KamLAND data. This result implies an upper bound of  $\sin^2\theta_{13} < 0.059$  (95% CL) or  $\sin^2 2\theta_{13} < 0.22$  (95% CL). The normal neutrino mass ordering and  $CP$  invariance are assumed.

<sup>19</sup> ABE 11 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{32}^2$  fixed to  $2.4 \times 10^{-3}$  eV<sup>2</sup>, using global solar neutrino data including Super-Kamiokande, SNO, Borexino (ARPESELLA 08A), Homestake, and GALLEX/GNO data. The normal neutrino mass ordering is assumed.

<sup>20</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.4 \times 10^{-3}$  eV<sup>2</sup>,  $\theta_{23} = \pi/2$ ,  $\delta = 0$ , and the normal mass ordering. For other values of  $\delta$ , the 68% region spans from 0.03 to 0.25, and the 90% region from 0.02 to 0.32.

<sup>21</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.4 \times 10^{-3}$  eV<sup>2</sup>,  $\theta_{23} = \pi/2$ ,  $\delta = 0$ , and the inverted mass ordering. For other values of  $\delta$ , the 68% region spans from 0.04 to 0.30, and the 90% region from 0.02 to 0.39.

<sup>22</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.32 \times 10^{-3}$  eV<sup>2</sup>,  $\theta_{23} = \pi/2$ ,  $\delta = 0$ , and the normal mass ordering. For other values of  $\delta$ , the 68% region spans from 0.02 to 0.12, and the 90% region from 0 to 0.16.

<sup>23</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.32 \times 10^{-3}$  eV<sup>2</sup>,  $\theta_{23} = \pi/2$ ,  $\delta = 0$ , and the inverted mass ordering. For other values of  $\delta$ , the 68% region spans from 0.02 to 0.16, and the 90% region from 0 to 0.21.

<sup>24</sup> FOGLI 11 obtained this result from an analysis using the atmospheric, accelerator long baseline, CHOOZ, solar, and KamLAND data. Recently, MUELLER 11 suggested an average increase of about 3.5% in normalization of the reactor  $\overline{\nu}_e$  fluxes, and using these fluxes, the fitted result becomes  $0.10 \pm 0.03$ .

<sup>25</sup> GANDO 11 report  $\sin^2\theta_{13} = 0.020 \pm 0.016$ . This result was obtained with three-neutrino fit using the KamLAND + solar data.

<sup>26</sup> GANDO 11 report  $\sin^2\theta_{13} = 0.032 \pm 0.037$ . This result was obtained with three-neutrino fit using the KamLAND data only.

<sup>27</sup> This result corresponds to the limit of  $<0.12$  at 90% CL for  $\Delta m_{32}^2 = 2.43 \times 10^{-3}$  eV<sup>2</sup>,  $\theta_{23} = \pi/2$ , and  $\delta = 0$ . For other values of  $\delta$ , the 90% CL region spans from 0 to 0.16.

<sup>28</sup> This result corresponds to the limit of  $<0.20$  at 90% CL for  $\Delta m_{32}^2 = 2.43 \times 10^{-3}$  eV<sup>2</sup>,  $\theta_{23} = \pi/2$ , and  $\delta = 0$ . For other values of  $\delta$ , the 90% CL region spans from 0 to 0.21.

<sup>29</sup> AHARMIM 10 global solar neutrino data include SNO's low-energy-threshold analysis survival probability day/night curves, SNO Phase III integral rates (AHARMIM 08), Cl (CLEVELAND 98), SAGE (ABDURASHITOV 09), GALLEX/GNO (HAMPEL 99, ALTMANN 05), Borexino (ARPESELLA 08A), SK-I zenith (HOSAKA 06), and SK-II day/night spectra (CRAVENS 08).

<sup>30</sup> AHARMIM 10 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.3 \times 10^{-3}$  eV<sup>2</sup>, using global solar neutrino data and KamLAND data (ABE 08A).  $CP$  invariance is assumed. This result implies an upper bound of  $\sin^2\theta_{13} < 0.057$  (95% CL) or  $\sin^2 2\theta_{13} < 0.22$  (95% CL).

# Lepton Particle Listings

## Neutrino Mixing

- <sup>31</sup> AHARMIM 10 obtained this result by a three-neutrino oscillation analysis with the value of  $\Delta m_{31}^2$  fixed to  $2.3 \times 10^{-3} \text{ eV}^2$ , using global solar neutrino data.
- <sup>32</sup> WENDELL 10 obtained this result by a three-neutrino oscillation analysis with one mass scale dominance ( $\Delta m_{21}^2 = 0$ ) using the Super-Kamiokande-I+II+III atmospheric neutrino data, and updates the HOSAKA 06A result.
- <sup>33</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.43 \times 10^{-3} \text{ eV}^2$ ,  $\theta_{23} = \pi/2$ , and  $\delta = 0$ . For other values of  $\delta$ , the 68% CL region spans from 0.02 to 0.26.
- <sup>34</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.43 \times 10^{-3} \text{ eV}^2$ ,  $\theta_{23} = \pi/2$ , and  $\delta = 0$ . For other values of  $\delta$ , the 68% CL region spans from 0.04 to 0.34.
- <sup>35</sup> FOGLI 08 obtained this result from a global analysis of all neutrino oscillation data, that is, solar + KamLAND + atmospheric + accelerator long baseline + CHOOZ.
- <sup>36</sup> FOGLI 08 obtained this result from an analysis using the solar and KamLAND neutrino oscillation data.
- <sup>37</sup> FOGLI 08 obtained this result from an analysis using the atmospheric, accelerator long baseline, and CHOOZ neutrino oscillation data.
- <sup>38</sup> YAMAMOTO 06 searched for  $\nu_\mu \rightarrow \nu_e$  appearance. Assumes  $2 \sin^2(2\theta_{\mu e}) = \sin^2(2\theta_{13})$ . The quoted limit is for  $\Delta m_{32}^2 = 1.9 \times 10^{-3} \text{ eV}^2$ . That value of  $\Delta m_{32}^2$  is the one- $\sigma$  low value for AHN 06A. For the AHN 06A best fit value of  $2.8 \times 10^{-3} \text{ eV}^2$ , the  $\sin^2(2\theta_{13})$  limit is  $< 0.26$ . Supersedes AHN 04.
- <sup>39</sup> AHN 04 searched for  $\nu_\mu \rightarrow \nu_e$  appearance. Assuming  $2 \sin^2(2\theta_{\mu e}) = \sin^2(2\theta_{13})$ , a limit on  $\sin^2(2\theta_{\mu e})$  is converted to a limit on  $\sin^2(2\theta_{13})$ . The quoted limit is for  $\Delta m_{32}^2 = 1.9 \times 10^{-3} \text{ eV}^2$ . That value of  $\Delta m_{32}^2$  is the one- $\sigma$  low value for ALIU 05. For the ALIU 05 best fit value of  $2.8 \times 10^{-3} \text{ eV}^2$ , the  $\sin^2(2\theta_{13})$  limit is  $< 0.30$ .
- <sup>40</sup> The quoted limit is for  $\Delta m_{32}^2 = 1.9 \times 10^{-3} \text{ eV}^2$ . That value of  $\Delta m_{32}^2$  is the 1- $\sigma$  low value for ALIU 05. For the ALIU 05 best fit value of  $2.8 \times 10^{-3} \text{ eV}^2$ , the  $\sin^2 2\theta_{13}$  limit is  $< 0.19$ . In this range, the  $\theta_{13}$  limit is larger for lower values of  $\Delta m_{32}^2$ , and smaller for higher values of  $\Delta m_{32}^2$ .
- <sup>41</sup> The quoted limit is for  $\Delta m_{32}^2 = 1.9 \times 10^{-3} \text{ eV}^2$ . That value of  $\Delta m_{32}^2$  is the 1- $\sigma$  low value for ALIU 05. For the ALIU 05 best fit value of  $2.8 \times 10^{-3} \text{ eV}^2$ , the  $\sin^2 2\theta_{13}$  limit is  $< 0.23$ .
- <sup>42</sup> The quoted limit is for  $\Delta m_{32}^2 = 2.43 \times 10^{-3} \text{ eV}^2$ . That value of  $\Delta m_{32}^2$  is the central value for ADAMSON 08. For the ADAMSON 08 1- $\sigma$  low value of  $2.30 \times 10^{-3} \text{ eV}^2$ , the  $\sin^2 2\theta_{13}$  limit is  $< 0.16$ . See also APOLLONIO 03 for a detailed description of the experiment.

### CP violating phase

#### $\delta$ , CP violating phase

Measurements of  $\delta$  come from atmospheric and accelerator experiments looking at  $\nu_e$  appearance. We encode values between 0 and  $2\pi$ , though it is equivalent to use  $-\pi$  to  $\pi$ .

VALUE ( $\pi$ rad)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0 to 0.15, 0.83 to 2	90	ABE 15D	T2K	Normal mass hierarchy
1.09 to 1.92	90	ABE 15D	T2K	Inverted mass hierarchy
0.05 to 1.2	90	<sup>1</sup> ADAMSON 14	MINS	Normal mass hierarchy
$1.34^{+0.64}_{-0.38}$		FORERO 14	FIT	Normal mass hierarchy
$1.48^{+0.34}_{-0.32}$		FORERO 14	FIT	Inverted mass hierarchy
$1.70^{+0.22}_{-0.39}$		<sup>2</sup> GONZALEZ-G..14	FIT	Normal mass hierarchy; global fit
$1.41^{+0.35}_{-0.34}$		<sup>2</sup> GONZALEZ-G..14	FIT	Inverted mass hierarchy; global fit
0 to 1.5 or 1.9 to 2	90	<sup>3</sup> ADAMSON 13A	MINS	Normal mass hierarchy

- <sup>1</sup> Based on three-flavor formalism and  $\theta_{23} > \pi/4$ . Likelihood as a function of  $\delta$  is also shown for the other three combinations of hierarchy and  $\theta_{23}$  quadrant; all values of  $\delta$  are allowed at 90% C.L.
- <sup>2</sup> GONZALEZ-GARCIA 14 result comes from a frequentist global fit. The corresponding Bayesian global fit to the same data results are reported in BERGSTROM 15 as 68% CL intervals of 1.24–1.94 for normal and 1.15–1.77 for inverted mass ordering.
- <sup>3</sup> Based on  $\nu_e$  appearance in MINOS and the calculated  $\sin^2(2\theta_{23}) = 0.957$ ,  $\theta_{23} > \pi/4$ , and normal mass hierarchy. Likelihood as a function of  $\delta$  is also shown for the other three combinations of hierarchy and  $\theta_{23}$  quadrant; all values of  $\delta$  are allowed at 90% C.L.

### (C) Other neutrino mixing results

The LSND collaboration reported in AGUILAR 01 a signal which is consistent with  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations. In a three neutrino framework, this would be a measurement of  $\theta_{12}$  and  $\Delta m_{21}^2$ . This does not appear to be consistent with most of the other neutrino data. The MiniBooNE experiment, reported in AGUILAR-AREVALO 07, does a two-neutrino analysis which, assuming  $CP$  conservation, rules out AGUILAR 01. However, the MiniBooNE antineutrino data reported in AGUILAR-AREVALO 13A are consistent with the signal reported in AGUILAR 01. The following listings include results which might be relevant towards understanding these observations. They include searches for  $\nu_\mu \rightarrow \nu_e$ ,  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ , sterile neutrino oscillations, and  $CPT$  violation.

### $\Delta(m^2)$ for $\sin^2(2\theta) = 1$ ( $\nu_\mu \rightarrow \nu_e$ )

VALUE (eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.015 to 0.050	90	<sup>1</sup> AGUILAR-AR...13A	MBOO	MiniBooNE
$< 0.34$	90	<sup>2</sup> MAHN 12	MBOO	MiniBooNE/SciBooNE
$< 0.034$	90	AGUILAR-AR...07	MBOO	MiniBooNE
$< 0.0008$	90	AHN 04	K2K	Water Cherenkov
$< 0.4$	90	ASTIER 03	NOMD	CERN SPS
$< 2.4$	90	AVVAKUMOV 02	NTEV	NUTEV FNAL
0.03 to 0.3	95	<sup>3</sup> AGUILAR 01	LSND	$\nu_\mu \rightarrow \nu_e$ osc.prob.
		<sup>4</sup> ATHANASSO...98	LSND	$\nu_\mu \rightarrow \nu_e$
		<sup>5</sup> LOVERRE 96	CHARM/CDHS	
$< 2.3$	90	VILAIN 94c	CHM2	CERN SPS
$< 0.9$	90	ANGELINI 86	HLBC	BEBC CERN PS

- <sup>1</sup> Based on  $\nu_\mu \rightarrow \nu_e$  appearance of  $162.0 \pm 47.8$  events; marginally compatible with two neutrino oscillations. The best fit value is  $\Delta m^2 = 3.14 \text{ eV}^2$ .
- <sup>2</sup> MAHN 12 is a combined spectral fit of MiniBooNE and SciBooNE neutrino data with the range of  $\Delta m^2$  up to  $25 \text{ eV}^2$ . The best limit is 0.04 at  $7 \text{ eV}^2$ .
- <sup>3</sup> AGUILAR 01 is the final analysis of the LSND full data set. Search is made for the  $\nu_\mu \rightarrow \nu_e$  oscillations using  $\nu_\mu$  from  $\pi^+$  decay in flight by observing beam-on electron events from  $\nu_e C \rightarrow e^- X$ . Present analysis results in  $8.1 \pm 12.2 \pm 1.7$  excess events in the  $60 < E_e < 200 \text{ MeV}$  energy range, corresponding to oscillation probability of  $0.10 \pm 0.16 \pm 0.04\%$ . This is consistent, though less significant, with the previous result of ATHANASSOPOULOS 98, which it supersedes. The present analysis uses selection criteria developed for the decay at rest region, and is less effective in removing the background above  $60 \text{ MeV}$  than ATHANASSOPOULOS 98.
- <sup>4</sup> ATHANASSOPOULOS 98 is a search for the  $\nu_\mu \rightarrow \nu_e$  oscillations using  $\nu_\mu$  from  $\pi^+$  decay in flight. The 40 observed beam-on electron events are consistent with  $\nu_e C \rightarrow e^- X$ ; the expected background is  $21.9 \pm 2.1$ . Authors interpret this excess as evidence for an oscillation signal corresponding to oscillations with probability  $(0.26 \pm 0.10 \pm 0.05)\%$ . Although the significance is only  $2.3\sigma$ , this measurement is an important and consistent cross check of ATHANASSOPOULOS 96 who reported evidence for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations from  $\mu^+$  decay at rest. See also ATHANASSOPOULOS 98b.
- <sup>5</sup> LOVERRE 96 uses the charged-current to neutral-current ratio from the combined CHARM (ALLABY 86) and CDHS (ABRAMOWICZ 86) data from 1986.

### $\sin^2(2\theta)$ for "Large" $\Delta(m^2)$ ( $\nu_\mu \rightarrow \nu_e$ )

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 7.2$	90	AGAFONOVA 13	OPER	$\Delta(m^2) > 0.1 \text{ eV}^2$
0.8 to 3	90	<sup>1</sup> AGUILAR-AR...13A	MBOO	MiniBooNE
$< 11$	90	<sup>2</sup> ANTONELLO 13	ICAR	$\nu_\mu \rightarrow \nu_e$
$< 6.8$	90	<sup>3</sup> ANTONELLO 13A	ICAR	$\nu_\mu \rightarrow \nu_e$
$< 100$	90	<sup>4</sup> MAHN 12	MBOO	MiniBooNE/SciBooNE
$< 1.8$	90	<sup>5</sup> AGUILAR-AR...07	MBOO	MiniBooNE
$< 110$	90	<sup>6</sup> AHN 04	K2K	Water Cherenkov
$< 1.4$	90	ASTIER 03	NOMD	CERN SPS
$< 1.6$	90	AVVAKUMOV 02	NTEV	NUTEV FNAL
0.5 to 30	95	<sup>7</sup> AGUILAR 01	LSND	$\nu_\mu \rightarrow \nu_e$ osc.prob.
		<sup>8</sup> ATHANASSO...98	LSND	$\nu_\mu \rightarrow \nu_e$
		<sup>9</sup> LOVERRE 96	CHARM/CDHS	
$< 3.0$	90	VILAIN 94c	CHM2	CERN SPS
$< 9.4$	90	VILAIN 94c	CHM2	CERN SPS
$< 5.6$	90	<sup>10</sup> VILAIN 94c	CHM2	CERN SPS

- <sup>1</sup> Based on  $\nu_\mu \rightarrow \nu_e$  appearance of  $162.0 \pm 47.8$  events; marginally compatible with two neutrino oscillations. The best fit value is  $\sin^2 2\theta = 0.002$ .
- <sup>2</sup> ANTONELLO 13 use the ICARUS T600 detector at LNGS and  $\sim 20 \text{ GeV}$  beam of  $\nu_\mu$  from CERN 730 km away to search for an excess of  $\nu_e$  events. Two events are found with  $3.7 \pm 0.6$  expected from conventional sources. This result excludes some parts of the parameter space expected by LSND. Superseded by ANTONELLO 13A.
- <sup>3</sup> Based on four events with a background of  $6.4 \pm 0.9$  from conventional sources with an average energy of  $20 \text{ GeV}$  and  $730 \text{ km}$  from the source of  $\nu_\mu$ .
- <sup>4</sup> MAHN 12 is a combined fit of MiniBooNE and SciBooNE neutrino data.
- <sup>5</sup> The limit is  $\sin^2 2\theta < 0.9 \times 10^{-3}$  at  $\Delta m^2 = 2 \text{ eV}^2$ . That value of  $\Delta m^2$  corresponds to the smallest mixing angle consistent with the reported signal from LSND in AGUILAR 01.
- <sup>6</sup> The limit becomes  $\sin^2 2\theta < 0.15$  at  $\Delta m^2 = 2.8 \times 10^{-3} \text{ eV}^2$ , the best-fit value of the  $\nu_\mu$  disappearance analysis in K2K.
- <sup>7</sup> AGUILAR 01 is the final analysis of the LSND full data set of the search for the  $\nu_\mu \rightarrow \nu_e$  oscillations. See footnote in preceding table for further details.
- <sup>8</sup> ATHANASSOPOULOS 98 report  $(0.26 \pm 0.10 \pm 0.05)\%$  for the oscillation probability; the value of  $\sin^2 2\theta$  for large  $\Delta m^2$  is deduced from this probability. See footnote in preceding table for further details, and see the paper for a plot showing allowed regions. If effect is due to oscillation, it is most likely to be intermediate  $\sin^2 2\theta$  and  $\Delta m^2$ . See also ATHANASSOPOULOS 98b.
- <sup>9</sup> LOVERRE 96 uses the charged-current to neutral-current ratio from the combined CHARM (ALLABY 86) and CDHS (ABRAMOWICZ 86) data from 1986.
- <sup>10</sup> VILAIN 94c limit derived by combining the  $\nu_\mu$  and  $\bar{\nu}_\mu$  data assuming  $CP$  conservation.



See key on page 601

# Lepton Particle Listings

## Neutrino Mixing

### $\Delta(m^2)$ for $\sin^2(2\theta) = 1$ ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ )

VALUE (eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.023 to 0.060	90	<sup>1</sup> AGUILAR-AR...13A	MBOO	MiniBooNE
<0.16	90	<sup>2</sup> CHENG 12	MBOO	MiniBooNE/SciBooNE
0.03–0.09	90	<sup>3</sup> AGUILAR-AR...10	MBOO	$E_\nu > 475$ MeV
0.03–0.07	90	<sup>4</sup> AGUILAR-AR...10	MBOO	$E_\nu > 200$ MeV
<0.06	90	AGUILAR-AR...09B	MBOO	MiniBooNE
<0.055	90	<sup>5</sup> ARMBRUSTER02	KAR2	Liquid Sci. calor.
<2.6	90	AVVAKUMOV 02	NTEV	NUTEV FNAL
0.03–0.05		<sup>6</sup> AGUILAR 01	LSND	LAMPF
0.05–0.08	90	<sup>7</sup> ATHANASSO...96	LSND	LAMPF
0.048–0.090	80	<sup>8</sup> ATHANASSO...95		
<0.07	90	<sup>9</sup> HILL 95		
<0.9	90	VILAIN 94c	CHM2	CERN SPS
<0.14	90	<sup>10</sup> FREEDMAN 93	CNTR	LAMPF

<sup>1</sup> Based on  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance of  $78.4 \pm 28.5$  events. The best fit values are  $\Delta m^2 = 0.043$  eV<sup>2</sup> and  $\sin^2 2\theta = 0.88$ .

<sup>2</sup> CHENG 12 is a combined fit of MiniBooNE and SciBooNE antineutrino data.

<sup>3</sup> This value is for a two neutrino oscillation analysis for excess antineutrino events with  $E_\nu > 475$  MeV. The best fit is at 0.07. The allowed region is consistent with LSND reported by AGUILAR 01. Supersedes AGUILAR-AREVALO 09B.

<sup>4</sup> This value is for a two neutrino oscillation analysis for excess antineutrino events with  $E_\nu > 200$  MeV with subtraction of the expected 12 events low energy excess seen in the neutrino component of the beam. The best fit value is 0.007 for  $\Delta(m^2) = 4.4$  eV<sup>2</sup>.

<sup>5</sup> ARMBRUSTER 02 is the final analysis of the KARMEN 2 data for 17.7 m distance from the ISIS stopped pion and muon neutrino source. It is a search for  $\bar{\nu}_e$ , detected by the inverse  $\beta$ -decay reaction on protons and <sup>12</sup>C. 15 candidate events are observed, and  $15.8 \pm 0.5$  background events are expected, hence no oscillation signal is detected. The results exclude large regions of the parameter area favored by the LSND experiment.

<sup>6</sup> AGUILAR 01 is the final analysis of the LSND full data set. It is a search for  $\bar{\nu}_e$  30 m from LAMPF beam stop. Neutrinos originate mainly from  $\pi^+$  decay at rest.  $\bar{\nu}_e$  are detected through  $\bar{\nu}_e p \rightarrow e^+ n$  ( $20 < E_{e^+} < 60$  MeV) in delayed coincidence with  $np \rightarrow d\gamma$ . Authors observe  $87.9 \pm 22.4 \pm 6.0$  total excess events. The observation is attributed to  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations with the oscillation probability of  $0.264 \pm 0.067 \pm 0.045\%$ , consistent with the previously published result. Taking into account all constraints, the most favored allowed region of oscillation parameters is a band of  $\Delta(m^2)$  from 0.2–2.0 eV<sup>2</sup>. Supersedes ATHANASSOPOULOS 95, ATHANASSOPOULOS 96, and ATHANASSOPOULOS 98.

<sup>7</sup> ATHANASSOPOULOS 96 is a search for  $\bar{\nu}_e$  30 m from LAMPF beam stop. Neutrinos originate mainly from  $\pi^+$  decay at rest.  $\bar{\nu}_e$  could come from either  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  or  $\nu_e \rightarrow \bar{\nu}_e$ ; our entry assumes the first interpretation. They are detected through  $\bar{\nu}_e p \rightarrow e^+ n$  ( $20 \text{ MeV} < E_{e^+} < 60 \text{ MeV}$ ) in delayed coincidence with  $np \rightarrow d\gamma$ . Authors observe  $51 \pm 20 \pm 8$  total excess events over an estimated background  $12.5 \pm 2.9$ . ATHANASSOPOULOS 96B is a shorter version of this paper.

<sup>8</sup> ATHANASSOPOULOS 95 error corresponds to the  $1.6\sigma$  band in the plot. The expected background is  $2.7 \pm 0.4$  events. Corresponds to an oscillation probability of  $(0.34^{+0.20}_{-0.18} \pm 0.07)\%$ . For a different interpretation, see HILL 95. Replaced by ATHANASSOPOULOS 96.

<sup>9</sup> HILL 95 is a report by one member of the LSND Collaboration, reporting a different conclusion from the analysis of the data of this experiment (see ATHANASSOPOULOS 95). Contrary to the rest of the LSND Collaboration, Hill finds no evidence for the neutrino oscillation  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  and obtains only upper limits.

<sup>10</sup> FREEDMAN 93 is a search at LAMPF for  $\bar{\nu}_e$  generated from any of the three neutrino types  $\nu_\mu$ ,  $\bar{\nu}_\mu$ , and  $\nu_e$  which come from the beam stop. The  $\bar{\nu}_e$ 's would be detected by the reaction  $\bar{\nu}_e p \rightarrow e^+ n$ . FREEDMAN 93 replaces DURKIN 88.

### $\sin^2(2\theta)$ for "Large" $\Delta(m^2)$ ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ )

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<640	90	<sup>1</sup> ANTONELLO 13A	ICAR	$\bar{\nu}_e$ appearance
<150	90	<sup>2</sup> CHENG 12	MBOO	MiniBooNE/SciBooNE
0.4–9.0	99	<sup>3</sup> AGUILAR-AR...10	MBOO	$E_\nu > 475$ MeV
0.4–9.0	99	<sup>4</sup> AGUILAR-AR...10	MBOO	$E_\nu > 200$ MeV
< 3.3	90	<sup>5</sup> AGUILAR-AR...09B	MBOO	MiniBooNE
< 1.7	90	<sup>6</sup> ARMBRUSTER02	KAR2	Liquid Sci. calor.
< 1.1	90	AVVAKUMOV 02	NTEV	NUTEV FNAL
5.3±1.3±9.0		<sup>7</sup> AGUILAR 01	LSND	LAMPF
6.2±2.4±1.0		<sup>8</sup> ATHANASSO...96	LSND	LAMPF
3–12	80	<sup>9</sup> ATHANASSO...95		
< 6	90	<sup>10</sup> HILL 95		

<sup>1</sup> ANTONELLO 13A obtained the limit by assuming  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation from the  $\sim 2\%$  of  $\bar{\nu}_\mu$  evnsts contamination in the CNGS beam.

<sup>2</sup> CHENG 12 is a combined fit of MiniBooNE and SciBooNE antineutrino data.

<sup>3</sup> This value is for a two neutrino oscillation analysis for excess antineutrino events with  $E_\nu > 475$  MeV. At 90% CL there is no solution at high  $\Delta(m^2)$ . The best fit is at maximal mixing. The allowed region is consistent with LSND reported by AGUILAR 01. Supersedes AGUILAR-AREVALO 09B.

<sup>4</sup> This value is for a two neutrino oscillation analysis for excess antineutrino events with  $E_\nu > 200$  MeV with subtraction of the expected 12 events low energy excess seen in the

neutrino component of the beam. At 90% CL there is no solution at high  $\Delta(m^2)$ . The best fit value is 0.007 for  $\Delta(m^2) = 4.4$  eV<sup>2</sup>.

<sup>5</sup> This result is inconclusive with respect to small amplitude mixing suggested by LSND.

<sup>6</sup> ARMBRUSTER 02 is the final analysis of the KARMEN 2 data. See footnote in the preceding table for further details, and the paper for the exclusion plot.

<sup>7</sup> AGUILAR 01 is the final analysis of the LSND full data set. The deduced oscillation probability is  $0.264 \pm 0.067 \pm 0.045\%$ ; the value of  $\sin^2 2\theta$  for large  $\Delta(m^2)$  is twice this probability (although these values are excluded by other constraints). See footnote in preceding table for further details, and the paper for a plot showing allowed regions. Supersedes ATHANASSOPOULOS 95, ATHANASSOPOULOS 96, and ATHANASSOPOULOS 98.

<sup>8</sup> ATHANASSOPOULOS 96 reports  $(0.31 \pm 0.12 \pm 0.05)\%$  for the oscillation probability; the value of  $\sin^2 2\theta$  for large  $\Delta(m^2)$  should be twice this probability. See footnote in preceding table for further details, and see the paper for a plot showing allowed regions.

<sup>9</sup> ATHANASSOPOULOS 95 error corresponds to the  $1.6\sigma$  band in the plot. The expected background is  $2.7 \pm 0.4$  events. Corresponds to an oscillation probability of  $(0.34^{+0.20}_{-0.18} \pm 0.07)\%$ . For a different interpretation, see HILL 95. Replaced by ATHANASSOPOULOS 96.

<sup>10</sup> HILL 95 is a report by one member of the LSND Collaboration, reporting a different conclusion from the analysis of the data of this experiment (see ATHANASSOPOULOS 95). Contrary to the rest of the LSND Collaboration, Hill finds no evidence for the neutrino oscillation  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  and obtains only upper limits.

### $\Delta(m^2)$ for $\sin^2(2\theta) = 1$ ( $\nu_\mu(\bar{\nu}_\mu) \rightarrow \nu_e(\bar{\nu}_e)$ )

VALUE (eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<0.075	90	BORODOV... 92	CNTR	BNL E776
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.6	90	<sup>1</sup> ROMOSAN 97	CCFR	FNAL

<sup>1</sup> ROMOSAN 97 uses wideband beam with a 0.5 km decay region.

### $\sin^2(2\theta)$ for "Large" $\Delta(m^2)$ ( $\nu_\mu(\bar{\nu}_\mu) \rightarrow \nu_e(\bar{\nu}_e)$ )

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.8	90	<sup>1</sup> ROMOSAN 97	CCFR	FNAL
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.8	90	<sup>2</sup> MCFARLAND 95	CCFR	FNAL
<3	90	BORODOV... 92	CNTR	BNL E776

<sup>1</sup> ROMOSAN 97 uses wideband beam with a 0.5 km decay region.

<sup>2</sup> MCFARLAND 95 state that "This result is the most stringent to date for  $250 < \Delta(m^2) < 450$  eV<sup>2</sup> and also excludes at 90%CL much of the high  $\Delta(m^2)$  region favored by the recent LSND observation." See ATHANASSOPOULOS 95 and ATHANASSOPOULOS 96.

### $\Delta(m^2)$ for $\sin^2(2\theta) = 1$ ( $\bar{\nu}_e \nrightarrow \bar{\nu}_e$ )

VALUE (eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.01	90	<sup>1</sup> ACHKAR 95	CNTR	Bugey reactor
<sup>1</sup> ACHKAR 95 bound is for $L=15, 40$ , and 95 m.				

### $\sin^2(2\theta)$ for "Large" $\Delta(m^2)$ ( $\bar{\nu}_e \nrightarrow \bar{\nu}_e$ )

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.02	90	<sup>1</sup> ACHKAR 95	CNTR	For $\Delta(m^2) = 0.6$ eV <sup>2</sup>
<sup>1</sup> ACHKAR 95 bound is from data for $L=15, 40$ , and 95 m distance from the Bugey reactor.				

### Sterile neutrino limits from atmospheric neutrino studies

### $\Delta(m^2)$ for $\sin^2(2\theta) = 1$ ( $\nu_\mu \rightarrow \nu_s$ )

$\nu_s$  means  $\nu_\tau$  or any sterile (noninteracting)  $\nu$ .

VALUE ( $10^{-5}$ eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3000 (or <550)	90	<sup>1</sup> OYAMA 89	KAMI	Water Cherenkov
<4.2 (or >54).	90	BIONTA 88	IMB	Flux has $\nu_\mu, \bar{\nu}_\mu, \nu_e$ , and $\bar{\nu}_e$

<sup>1</sup> OYAMA 89 gives a range of limits, depending on assumptions in their analysis. They argue that the region  $\Delta(m^2) = (100\text{--}1000) \times 10^{-5}$  eV<sup>2</sup> is not ruled out by any data for large mixing.

### Search for $\nu_\mu \rightarrow \nu_s$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<sup>1</sup> AMBROSIO 01	MCRO	matter effects	
<sup>2</sup> FUKUDA 00	SKAM	neutral currents + matter effects	

# Lepton Particle Listings

## Neutrino Mixing

<sup>1</sup> AMBROSIO 01 tested the pure 2-flavor  $\nu_\mu \rightarrow \nu_s$  hypothesis using matter effects which change the shape of the zenith-angle distribution of upward through-going muons. With maximum mixing and  $\Delta(m^2)$  around  $0.0024\text{ eV}^2$ , the  $\nu_\mu \rightarrow \nu_s$  oscillation is disfavored with 99% confidence level with respect to the  $\nu_\mu \rightarrow \nu_\tau$  hypothesis.

<sup>2</sup> FUKUDA 00 tested the pure 2-flavor  $\nu_\mu \rightarrow \nu_s$  hypothesis using three complementary atmospheric-neutrino data samples. With this hypothesis, zenith-angle distributions are expected to show characteristic behavior due to neutral currents and matter effects. In the  $(\Delta(m^2))$  and  $\sin^2 2\theta$  region preferred by the Super-Kamiokande data, the  $\nu_\mu \rightarrow \nu_s$  hypothesis is rejected at the 99% confidence level, while the  $\nu_\mu \rightarrow \nu_\tau$  hypothesis consistently fits all of the data sample.

### CPT tests

$\langle \Delta m_{21}^2 - \Delta m_{21}^2 \rangle$

VALUE (10 <sup>-4</sup> eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.1 99.7 <sup>1</sup> DEGOUVEA 05 FIT solar vs. reactor

<sup>1</sup> DEGOUVEA 05 obtained this bound at the  $3\sigma$  CL from the KamLAND (ARAKI 05) and solar neutrino data.

$\langle \Delta m_{32}^2 - \Delta m_{32}^2 \rangle$

VALUE (10 <sup>-3</sup> eV <sup>2</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.6  $\pm$  2.4 90 <sup>1</sup> ADAMSON 12B MINS MINOS atmospheric

<sup>1</sup> The quoted result is the single-parameter 90% C.L. interval determined from the 90% C.L. contour in the  $(\Delta m^2, \Delta m^2)$  plane, which is obtained by minimizing the four parameter log-likelihood function with respect to the other oscillation parameters.

### REFERENCES FOR Neutrino Mixing

AARTSEN 15A PR D91 072004 M. G. Aartsen (IceCube Collab.)  
ABE 15D PR D91 072010 K. Abe et al. (T2K Collab.)  
AN 15 PRL 115 111802 F.P. An et al. (Daya Bay Collab.)  
BERGSTROM 15 JHEP 1509 200 J. Bergstrom et al. (BARC, STON, MADU+)  
GANDO 15 PR C92 055808 A. Gando et al. (KamLAND Collab.)  
ABE 14 PRL 112 181801 K. Abe et al. (T2K Collab.)  
Also PR D91 072010 Y. Abe et al. (T2K Collab.)  
ABE 14A PRL 112 181801 K. Abe et al. (T2K Collab.)  
ABE 14B PR D89 092003 K. Abe et al. (T2K Collab.)  
ABE 14C PRL 112 061802 K. Abe et al. (T2K Collab.)  
ABE 14H JHEP 1410 086 Y. Abe et al. (Double Chooz Collab.)  
Also JHEP 1502 074 (errat.) Y. Abe et al. (Double Chooz Collab.)  
ADAMSON 14 PRL 112 191801 P. Adamson et al. (MINOS Collab.)  
AN 14 PRL 112 061801 F.P. An et al. (Daya Bay Collab.)  
AN 14B PR D90 071101 F.P. An et al. (Daya Bay Collab.)  
BELLINI 14A NAT 512 383 G. Bellini et al. (Borexino Collab.)  
FORERO 14 PR D90 093006 D. V. Forero, M. Tortola, J. W. F. Valle  
GONZALEZ-G... 14 JHEP 1411 052 M.C. Gonzalez-Garcia, M. Maltoni, T. Schwetz  
PDG 14 CPC 38 070001 K. Olive et al. (PDG Collab.)  
RENSHAW 14 PRL 112 091805 A. Renshaw et al. (Super-Kamiokande Collab.)  
AARTSEN 13B PRL 111 081801 M. G. Aartsen et al. (IceCube Collab.)  
ABE 13C PL B723 66 Y. Abe et al. (Double Chooz Collab.)  
ABE 13E PR D88 032002 K. Abe et al. (T2K Collab.)  
ABE 13G PRL 111 211803 K. Abe et al. (T2K Collab.)  
ADAMSON 13A PRL 110 171801 P. Adamson et al. (MINOS Collab.)  
ADAMSON 13B PRL 110 251801 P. Adamson et al. (MINOS Collab.)  
AGAFONOVA 13 JHEP 1307 004 N. Agafonova et al. (OPERA Collab.)  
AGUILAR-AR... 13A PRL 110 161801 A.A. Aguilar-Arevalo et al. (MiniBooNE Collab.)  
AHARMIM 13 PR C88 025501 B. Aharmim et al. (SNO Collab.)  
AN 13 CPC 37 011001 F.P. An et al. (Daya Bay Collab.)  
Also CPC 37 011001 (errat.) F.P. An et al. (Daya Bay Collab.)  
ANTONELLO 13 EPJ C73 2345 M. Antonello et al. (ICARUS Collab.)  
ANTONELLO 13A EPJ C73 2599 M. Antonello et al. (ICARUS Collab.)  
GANDO 13 PR D88 033001 A. Gando et al. (KamLAND Collab.)  
ABE 12 PRL 108 131801 Y. Abe et al. (Double Chooz Collab.)  
ABE 12A PR D85 031103 K. Abe et al. (T2K Collab.)  
ABE 12B PR D86 052008 Y. Abe et al. (Double Chooz Collab.)  
ADAMSON 12 PRL 108 191801 P. Adamson et al. (MINOS Collab.)  
ADAMSON 12B PR D86 052007 P. Adamson et al. (MINOS Collab.)  
ADRIAN-MAR... 12 PL B714 224 S. Adrian-Martinez et al. (ANTARES Collab.)  
JHN 12 PRL 108 171802 J.K. Ahn et al. (RENO Collab.)  
AN 12 PRL 108 171803 F.P. An et al. (Daya Bay Collab.)  
Also CPC 37 011001 (errat.) F.P. An et al. (Daya Bay Collab.)  
BELLINI 12A PRL 108 051302 G. Bellini et al. (Borexino Collab.)  
CHENG 12 PR D86 052009 G. Cheng et al. (MiniBooNE/SciBooNE Collab.)  
MAHN 12 PR D85 032007 K.B.M. Mahn et al. (MiniBooNE/SciBooNE Collab.)  
ABE 11 PR D83 052010 K. Abe et al. (Super-Kamiokande Collab.)  
ABE 11A PRL 107 041801 K. Abe et al. (T2K Collab.)  
ABE 11B PR C84 035804 S. Abe et al. (KamLAND Collab.)  
ABE 11C PRL 107 241801 K. Abe et al. (Super-Kamiokande Collab.)  
ADAMSON 11 PRL 106 181801 P. Adamson et al. (MINOS Collab.)  
ADAMSON 11B PRL 107 021801 P. Adamson et al. (MINOS Collab.)  
ADAMSON 11C PR D84 071103 P. Adamson et al. (MINOS Collab.)  
ADAMSON 11D PRL 107 181802 P. Adamson et al. (MINOS Collab.)  
BELLINI 11 PL B696 191 G. Bellini et al. (Borexino Collab.)  
BELLINI 11A PRL 107 141302 G. Bellini et al. (Borexino Collab.)  
FOGLI 11 PR D84 053007 G.L. Fogli et al. (KamLAND Collab.)  
GANDO 11 PR D83 022002 A. Gando et al. (KamLAND Collab.)  
MUELLER 11 PR C83 054615 Th.A. Mueller et al.  
SERENELLI 11 APJ 743 24 A.M. Serenelli, W.C. Haxton, C. Pena-Garay  
ADAMSON 10A PR D82 051102 P. Adamson et al. (MINOS Collab.)  
AGUILAR-AR... 10 PRL 105 181801 A.A. Aguilar-Arevalo et al. (MiniBooNE Collab.)  
AHARMIM 10 PR C81 055504 B. Aharmim et al. (SNO Collab.)  
BELLINI 10A PR D82 033006 G. Bellini et al. (Borexino Collab.)  
DENIZ 10 PR D81 072001 M. Deniz et al. (TEXONO Collab.)  
KAETHER 10 PL B685 47 F. Kaether et al.  
WENDELL 10 PR D81 092004 R. Wendell et al. (Super-Kamiokande Collab.)

ABDURASHIL... 09 PR C80 015807 J.N. Abdurashitov et al. (SAGE Collab.)  
ADAMSON 09 PRL 103 261802 P. Adamson et al. (MINOS Collab.)  
AGUILAR-AR... 09B PRL 103 111801 A.A. Aguilar-Arevalo et al. (MiniBooNE Collab.)  
ABE 08A PRL 100 221803 S. Abe et al. (KamLAND Collab.)  
Also PRL 101 119904 S. Abe et al. (KamLAND Collab.)  
ADAMSON 08 PR D77 072002 P. Adamson et al. (MINOS Collab.)  
ADAMSON 08A PRL 101 131802 P. Adamson et al. (MINOS Collab.)  
AHARMIM 08 PRL 101 111301 B. Aharmim et al. (SNO Collab.)  
Also PR C87 015502 B. Aharmim et al. (SNO Collab.)  
ARPESELLA 08A PRL 101 091302 C. Arpesella et al. (Borexino Collab.)  
CRAVENS 08 PR D78 032002 J.P. Cravens et al. (Super-Kamiokande Collab.)  
FOGLI 08 PRL 101 141801 G.L. Fogli, et al.  
ADAMSON 07 PR D75 092003 P. Adamson et al. (MINOS Collab.)  
AGUILAR-AR... 07 PRL 98 231801 A.A. Aguilar-Arevalo et al. (MiniBooNE Collab.)  
AHARMIM 07 PR C75 045502 B. Aharmim et al. (SNO Collab.)  
ADAMSON 06 PR D73 072002 P. Adamson et al. (MINOS Collab.)  
AHN 06A PR D74 072003 M.H. Ahn et al. (K2K Collab.)  
BALATA 06 EPJ C47 21 M. Balata et al. (Borexino Collab.)  
HOSAKA 06 PR D73 122001 J. Hosaka et al. (Super-Kamiokande Collab.)  
HOSAKA 06A PR D74 032002 J. Hosaka et al. (Super-Kamiokande Collab.)  
MICHAEL 06 PRL 97 191801 D. Michael et al. (MINOS Collab.)  
WINTER 06 PR C73 025503 W.T. Winter et al.  
YAMAMOTO 06A PRL 96 181801 S. Yamamoto et al. (K2K Collab.)  
AHARMIM 05A PR C72 055502 B. Aharmim et al. (SNO Collab.)  
ALIU 05 PRL 94 081802 E. Aliu et al. (K2K Collab.)  
ALLISON 05 PR D72 052005 W.W.M. Allison et al. (SOUDAN-2 Collab.)  
ALTMANN 05 PL B616 174 M. Altmann et al. (GNO Collab.)  
ARAKI 05 PRL 94 081801 T. Araki et al. (KamLAND Collab.)  
ASHIE 05 PR D71 122005 Y. Ashie et al. (Super-Kamiokande Collab.)  
DEGOUVEA 05 PR D71 093002 A. de Gouvea, C. Pena-Garay  
AHARMIM 04 PR D70 090104 B. Aharmim et al. (SNO Collab.)  
AHMED 04A PRL 92 181301 S.N. Ahmed et al. (SNO Collab.)  
AHN 04 PRL 93 051801 M.H. Ahn et al. (K2K Collab.)  
AMBROSIO 04 EPJ C36 323 M. Ambrosio et al. (MACRO Collab.)  
ASHIE 04 PRL 93 101801 Y. Ashie et al. (Super-Kamiokande Collab.)  
EGUCHI 04 PRL 92 071301 K. Eguchi et al. (KamLAND Collab.)  
SMY 04 PR D69 011104 M.B. Smy et al. (Super-Kamiokande Collab.)  
AHN 03 PRL 90 041801 M.H. Ahn et al. (K2K Collab.)  
AMBROSIO 03 PL B566 35 M. Ambrosio et al. (MACRO Collab.)  
APOLLONIO 03 EPJ C27 331 M. Apollonio et al. (CHOOZ Collab.)  
ASTIER 03 PL B570 19 P. Astier et al. (NOMAD Collab.)  
EGUCHI 03 PR D69 021802 K. Eguchi et al. (KamLAND Collab.)  
GANDO 03 PRL 90 171302 Y. Gando et al. (Super-Kamiokande Collab.)  
IANNI 03 JP G29 2107 A. Ianni (INFN Gran Sasso)  
SANCHEZ 03 PR D68 113004 M. Sanchez et al. (Soudan 2 Collab.)  
ABDURASHIL... 02 JETP 95 181 J.N. Abdurashitov et al. (SAGE Collab.)  
Also Translated from ZETF 122 211.  
AHMAD 02 PRL 89 011301 Q.R. Ahmad et al. (SNO Collab.)  
AHMAD 02B PRL 89 011302 Q.R. Ahmad et al. (SNO Collab.)  
ARMBRUSTER 02 PR D65 112001 B. Armbruster et al. (KARMEN 2 Collab.)  
AVVAKUMOV 02 PRL 89 011804 S. Avvakumov et al. (NuTeV Collab.)  
FUKUDA 02 PL B539 179 S. Fukuda et al. (Super-Kamiokande Collab.)  
AGUILAR 01 PR D64 122007 A. Aguilar et al. (LSND Collab.)  
AHMAD 01 PRL 87 071301 Q.R. Ahmad et al. (SNO Collab.)  
AMBROSIO 01 PL B517 59 M. Ambrosio et al. (MACRO Collab.)  
BOEHM 01 PR D64 112001 F. Boehm et al.  
FUKUDA 01 PRL 86 5651 S. Fukuda et al. (Super-Kamiokande Collab.)  
AMBROSIO 00 PL B478 5 M. Ambrosio et al. (MACRO Collab.)  
BOEHM 00 PRL 84 3764 F. Boehm et al.  
FUKUDA 00 PRL 85 3999 S. Fukuda et al. (Super-Kamiokande Collab.)  
ALLISON 99 PL B449 137 W.W.M. Allison et al. (Soudan 2 Collab.)  
APOLLONIO 99 PL B466 415 M. Apollonio et al. (CHOOZ Collab.)  
Also PRL B472 434 (errat.) M. Apollonio et al. (CHOOZ Collab.)  
FUKUDA 99C PRL 82 2644 Y. Fukuda et al. (Super-Kamiokande Collab.)  
FUKUDA 99D PL B467 185 Y. Fukuda et al. (Super-Kamiokande Collab.)  
HAMPEL 99 PL B447 127 W. Hampel et al. (GALLEx Collab.)  
AMBROSIO 98 PL B434 451 M. Ambrosio et al. (MACRO Collab.)  
APOLLONIO 98 PL B420 397 M. Apollonio et al. (CHOOZ Collab.)  
ATHANASSO... 98 PRL 81 1774 C. Athanassopoulos et al. (LSND Collab.)  
ATHANASSO... 98B PR C58 2489 C. Athanassopoulos et al. (LSND Collab.)  
CLEVELAND 98 APJ 496 505 B.T. Cleveland et al. (Homestake Collab.)  
FELDMAN 98 PR D57 3873 G.J. Feldman, R.D. Cousins  
FUKUDA 98C PRL 81 1562 Y. Fukuda et al. (Super-Kamiokande Collab.)  
HATAKEYAMA 98 PRL 81 2016 S. Hatakeyama et al. (Kamiokande Collab.)  
CLARK 97 PRL 79 345 R. Clark et al. (IMB Collab.)  
ROMOSAN 97 PRL 78 2912 A. Romosan et al. (CCFR Collab.)  
AGLIETTA 96 JETPL 63 791 M. Aglietta et al. (LSD Collab.)  
Also Translated from ZETFP 63 753.  
ATHANASSO... 96 PR C54 2685 C. Athanassopoulos et al. (LSND Collab.)  
ATHANASSO... 96B PR 77 3082 C. Athanassopoulos et al. (LSND Collab.)  
FUKUDA 96 PRL 77 1683 Y. Fukuda et al. (Kamiokande Collab.)  
FUKUDA 96B PL B388 397 Y. Fukuda et al. (Kamiokande Collab.)  
GREENWOOD 96 PR D53 6054 Z.D. Greenwood et al. (UCI, SVR, SCUC)  
HAMPEL 96 PL B388 384 W. Hampel et al. (GALLEx Collab.)  
LOVERRE 96 PL B370 156 P.F. Loverre  
ACHKAR 95 NP B434 503 B. Achkar et al. (SING, SACL, CPPM, CDEF+)  
AHLEN 95 PR B357 481 S.P. Ahlen et al. (MACRO Collab.)  
ATHANASSO... 95 PRL 75 2650 C. Athanassopoulos et al. (LSND Collab.)  
DAUM 95 ZPHY C66 417 K. Daum et al. (FREJUS Collab.)  
HILL 95 PRL 75 2654 J.E. Hill (PENN)  
MCFARLAND 95 PRL 75 3933 K.S. McFarland et al. (CCFR Collab.)  
DECLAIS 94 PL B338 383 Y. Declais et al.  
FUKUDA 94 PL B335 237 Y. Fukuda et al. (Kamiokande Collab.)  
VILAIN 94C ZPHY C64 539 P. Vilain et al. (CHARM II Collab.)  
FREEDMAN 93 PR D47 811 S.J. Freedman et al. (LAMPF E645 Collab.)  
BECKER-SZ... 92B PR D46 3720 R.A. Becker-Szendy et al. (IMB Collab.)  
BEIER 92 PL B283 446 E.W. Beier et al. (KAM2 Collab.)  
Also PTRSL A346 63 E.W. Beier, E.D. Frank (PENN)  
BORODOV... 92 PRL 68 274 L. Borodovsky et al. (COLU, JHU, ILL)  
HIRATA 92 PRL B280 146 K.S. Hirata et al. (Kamiokande II Collab.)  
CASPER 91 PRL 66 2561 D. Casper et al. (IMB Collab.)  
HIRATA 91 PRL 66 9 ANGELINI 91 K.S. Hirata et al. (Kamiokande II Collab.)  
KUVSHIN... 91 JETPL 54 253 A.A. Kuvsinnikov et al. (KIAE)  
BERGER 90B PL B245 305 C. Berger et al. (FREJUS Collab.)  
HIRATA 90 PRL 65 1297 K.S. Hirata et al. (Kamiokande II Collab.)  
AGLIETTA 89 EPL 8 611 M. Aglietta et al. (FREJUS Collab.)  
DAVIS 89 ARNPS 39 467 R. Davis, A.K. Mann, L. Wolfenstein (BNL, PENN+)  
OYAMA 89 PR D39 1481 Y. Oyama et al. (Kamiokande II Collab.)  
BIONTA 88 PR D38 768 R.M. Bionta et al. (IMB Collab.)  
DURKIN 88 PRL 61 1811 L.S. Durkin et al. (OSU, ANL, CIT+)  
ABRAMOWICZ 86 PRL 57 298 H. Abramowicz et al. (CDHS Collab.)  
ALLABY 86 PL B177 446 J.V. Allaby et al. (CHARM Collab.)  
ANGELINI 84 PL B179 307 C. Angelini et al. (PISA, ATLAS, PAD+)  
VUILLEUMIER 82 PL 114B 298 J.L. Vuilleumier et al. (PISA, ATLAS, PAD+)  
BOLIV 81 SNJP 34 787 M.M. Boliev et al. (INRM)  
Also Translated from YAF 34 1418.  
KWON 81 PR D24 1097 H. Kwon et al. (CIT, ISNG, MUNI)  
BOEHM 80 PL 97B 310 F. Boehm et al. (ILLG, CIT, ISNG, MUNI)  
CROUCH 78 PR D18 2319 M.F. Crouch et al. (CASE, UCI, WITW)

See key on page 601

# Lepton Particle Listings

## Heavy Neutral Leptons, Searches for

### Heavy Neutral Leptons, Searches for

#### (A) Heavy Neutral Leptons

##### Stable Neutral Heavy Lepton MASS LIMITS

Note that LEP results in combination with REUSSER 91 exclude a fourth stable neutrino with  $m < 2400$  GeV.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>45.0	95	ABREU	92B	DLPH Dirac
>39.5	95	ABREU	92B	DLPH Majorana
>44.1	95	ALEXANDER	91F	OPAL Dirac
>37.2	95	ALEXANDER	91F	OPAL Majorana
none 3–100	90	SATO	91	KAM2 Kamiokande II
>42.8	95	<sup>1</sup> ADEVA	90S	L3 Dirac
>34.8	95	<sup>1</sup> ADEVA	90S	L3 Majorana
>42.7	95	DECAMP	90F	ALEP Dirac

<sup>1</sup>ADEVA 90s limits for the heavy neutrino apply if the mixing with the charged leptons satisfies  $|U_{1j}|^2 + |U_{2j}|^2 + |U_{3j}|^2 > 6.2 \times 10^{-8}$  at  $m_{L^0} = 20$  GeV and  $> 5.1 \times 10^{-10}$  for  $m_{L^0} = 40$  GeV.

##### Heavy Neutral Lepton MASS LIMITS

Limits apply only to heavy lepton type given in comment at right of data Listings.

See the “Quark and Lepton Compositeness, Searches for” Listings for limits on radiatively decaying excited neutral leptons, i.e.  $\nu^* \rightarrow \nu\gamma$ .

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>101.3	95	ACHARD	01B	L3 Dirac coupling to $e$
>101.5	95	ACHARD	01B	L3 Dirac coupling to $\mu$
> 90.3	95	ACHARD	01B	L3 Dirac coupling to $\tau$
> 89.5	95	ACHARD	01B	L3 Majorana coupling to $e$
> 90.7	95	ACHARD	01B	L3 Majorana coupling to $\mu$
> 80.5	95	ACHARD	01B	L3 Majorana coupling to $\tau$

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 76.0	95	ABBIENDI	00i	OPAL Majorana, coupling to $e$
> 88.0	95	ABBIENDI	00i	OPAL Dirac, coupling to $e$
> 76.0	95	ABBIENDI	00i	OPAL Majorana, coupling to $\mu$
> 88.1	95	ABBIENDI	00i	OPAL Dirac, coupling to $\mu$
> 53.8	95	ABBIENDI	00i	OPAL Majorana, coupling to $\tau$
> 71.1	95	ABBIENDI	00i	OPAL Dirac, coupling to $\tau$
> 76.5	95	ABREU	990	DLPH Dirac coupling to $e$
> 79.5	95	ABREU	990	DLPH Dirac coupling to $\mu$
> 60.5	95	ABREU	990	DLPH Dirac coupling to $\tau$
> 63	95	<sup>2,3</sup> BUSKULIC	96S	ALEP Dirac
> 54.3	95	<sup>2,4</sup> BUSKULIC	96S	ALEP Majorana

<sup>2</sup>BUSKULIC 96s requires the decay length of the heavy lepton to be  $< 1$  cm, limiting the square of the mixing angle  $|U_{ej}|^2$  to  $10^{-10}$ .

<sup>3</sup>BUSKULIC 96s limit for mixing with  $\tau$ . Mass is  $> 63.6$  GeV for mixing with  $e$  or  $\mu$ .

<sup>4</sup>BUSKULIC 96s limit for mixing with  $\tau$ . Mass is  $> 55.2$  GeV for mixing with  $e$  or  $\mu$ .

##### Astrophysical Limits on Neutrino MASS for $m_\nu > 1$ GeV

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
none 60–115		<sup>5</sup> FARGION	95	ASTR Dirac
none 9.2–2000		<sup>6</sup> GARCIA	95	COSM Nucleosynthesis
none 26–4700		<sup>6</sup> BECK	94	COSM Dirac
none 6 – hundreds		<sup>7,8</sup> MORI	92B	KAM2 Dirac neutrino
none 24 – hundreds		<sup>7,8</sup> MORI	92B	KAM2 Majorana neutrino
none 10–2400	90	<sup>9</sup> REUSSER	91	CNTR HPGe search
none 3–100	90	SATO	91	KAM2 Kamiokande II
		<sup>10</sup> ENQVIST	89	COSM
none 12–1400		<sup>6</sup> CALDWELL	88	COSM Dirac $\nu$
none 4–16	90	<sup>6,7</sup> OLIVE	88	COSM Dirac $\nu$
none 4–35	90	OLIVE	88	COSM Majorana $\nu$
>4.2 to 4.7		SREDNICKI	88	COSM Dirac $\nu$
>5.3 to 7.4		SREDNICKI	88	COSM Majorana $\nu$
none 20–1000	95	<sup>6</sup> AHLEN	87	COSM Dirac $\nu$
>4.1		GRIEST	87	COSM Dirac $\nu$

<sup>5</sup>FARGION 95 bound is sensitive to assumed  $\nu$  concentration in the Galaxy. See also KONOPLICH 94.

<sup>6</sup>These results assume that neutrinos make up dark matter in the galactic halo.

<sup>7</sup>Limits based on annihilations in the sun and are due to an absence of high energy neutrinos detected in underground experiments.

<sup>8</sup>MORI 92B results assume that neutrinos make up dark matter in the galactic halo. Limits based on annihilations in earth are also given.

<sup>9</sup>REUSSER 91 uses existing  $\beta\beta$  detector (see FISHER 89) to search for CDM Dirac neutrinos.

<sup>10</sup>ENQVIST 89 argue that there is no cosmological upper bound on heavy neutrinos.

#### (B) Other Bounds from Nuclear and Particle Decays

##### Limits on $|U_{ex}|^2$ as Function of $m_{\nu_x}$

##### Peak and kink search tests

Limits on  $|U_{ex}|^2$  as function of  $m_{\nu_j}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<1 $\times 10^{-7}$	90	<sup>11</sup> BRITTON	92B	CNTR 50 MeV $< m_{\nu_x} < 130$ MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5 $\times 10^{-6}$	90	DELEENER...	91	$m_{\nu_x} = 20$ MeV
<5 $\times 10^{-7}$	90	DELEENER...	91	$m_{\nu_x} = 40$ MeV
<3 $\times 10^{-7}$	90	DELEENER...	91	$m_{\nu_x} = 60$ MeV
<1 $\times 10^{-6}$	90	DELEENER...	91	$m_{\nu_x} = 80$ MeV
<1 $\times 10^{-6}$	90	DELEENER...	91	$m_{\nu_x} = 100$ MeV
<5 $\times 10^{-7}$	90	AZUELOS	86	CNTR $m_{\nu_x} = 60$ MeV
<2 $\times 10^{-7}$	90	AZUELOS	86	CNTR $m_{\nu_x} = 80$ MeV
<3 $\times 10^{-7}$	90	AZUELOS	86	CNTR $m_{\nu_x} = 100$ MeV
<1 $\times 10^{-6}$	90	AZUELOS	86	CNTR $m_{\nu_x} = 120$ MeV
<2 $\times 10^{-7}$	90	AZUELOS	86	CNTR $m_{\nu_x} = 130$ MeV
<1 $\times 10^{-4}$	90	<sup>12</sup> BRYMAN	83B	CNTR $m_{\nu_x} = 5$ MeV
<1.5 $\times 10^{-6}$	90	BRYMAN	83B	CNTR $m_{\nu_x} = 53$ MeV
<1 $\times 10^{-5}$	90	BRYMAN	83B	CNTR $m_{\nu_x} = 70$ MeV
<1 $\times 10^{-4}$	90	BRYMAN	83B	CNTR $m_{\nu_x} = 130$ MeV
<1 $\times 10^{-4}$	68	<sup>13</sup> SHROCK	81	THEO $m_{\nu_x} = 10$ MeV
<5 $\times 10^{-6}$	68	<sup>13</sup> SHROCK	81	THEO $m_{\nu_x} = 60$ MeV
<1 $\times 10^{-5}$	68	<sup>14</sup> SHROCK	80	THEO $m_{\nu_x} = 80$ MeV
<3 $\times 10^{-6}$	68	<sup>14</sup> SHROCK	80	THEO $m_{\nu_x} = 160$ MeV

<sup>11</sup>BRITTON 92B is from a search for additional peaks in the  $e^+$  spectrum from  $\pi^+ \rightarrow e^+ \nu_e$  decay at TRIUMF. See also BRITTON 92.

<sup>12</sup>BRYMAN 83B obtain upper limits from both direct peak search and analysis of  $B(\pi \rightarrow e\nu)/B(\pi \rightarrow \mu\nu)$ . Latter limits are not listed, except for this entry (i.e. — we list the most stringent limits for given mass).

<sup>13</sup>Analysis of  $(\pi^+ \rightarrow e^+ \nu_e)/(\pi^+ \rightarrow \mu^+ \nu_\mu)$  and  $(K^+ \rightarrow e^+ \nu_e)/(K^+ \rightarrow \mu^+ \nu_\mu)$  decay ratios.

<sup>14</sup>Analysis of  $(K^+ \rightarrow e^+ \nu_e)$  spectrum.

##### Kink search in nuclear $\beta$ decay

High-sensitivity follow-up experiments show that indications for a neutrino with mass 17 keV (Simpson, Hime, and others) were not valid. Accordingly, we no longer list the experiments by these authors and some others which made positive claims of 17 keV neutrino emission. Complete listings are given in the 1994 edition (Physical Review **D50** 1173 (1994)) and in the 1998 edition (The European Physical Journal **C3** 1 (1998)). We list below only the best limits on  $|U_{ex}|^2$  for each  $m_{\nu_x}$ . See WIETFELDT 96 for a comprehensive review.

VALUE (units $10^{-3}$ )	CL%	$m_{\nu_j}$ (keV)	ISOTOPE	METHOD	DOCUMENT ID
< 4–20	90	700–3500	<sup>38m</sup> K	Trap	<sup>15</sup> TRINCZEK 03
< 9–116	95	1–0.1	<sup>187</sup> Re	cryog.	<sup>16</sup> GALEAZZI 01
< 1	95	10–90	<sup>35</sup> S	Mag spect	<sup>17</sup> HOLZSCHUH 00
< 4	95	14–17	<sup>241</sup> Pu	Electrostatic spec	<sup>18</sup> DRAGOUN 99
< 1	95	4–30	<sup>63</sup> Ni	Mag spect	<sup>19</sup> HOLZSCHUH 99
< 10–40	90	370–640	<sup>37</sup> Ar	EC ion recoil	<sup>20</sup> HINDI 98
< 10	95	1	<sup>3</sup> H	SPEC	<sup>21</sup> HIDDEMANN 95
< 6	95	2	<sup>3</sup> H	SPEC	<sup>21</sup> HIDDEMANN 95
< 2	95	3	<sup>3</sup> H	SPEC	<sup>21</sup> HIDDEMANN 95
< 0.7	99	16.3–16.6	<sup>35</sup> S	Prop chamber	<sup>22</sup> KALBFLEISCH 93
< 2	95	13–40	<sup>3</sup> S	Si(Li)	<sup>23</sup> MORTARA 93
< 0.73	95	17	<sup>63</sup> Ni	Mag spect	<sup>24</sup> OHSHIMA 93
< 1.0	95	10–24	<sup>63</sup> Ni	Mag spect	<sup>25</sup> KAWAKAMI 92
< 0.9–2.5	90	1200–6800	<sup>20</sup> F	beta spectrum	<sup>26</sup> DEUTSCH 90
< 8	90	80	<sup>35</sup> S	Mag spect	<sup>27</sup> APALIKOV 85
< 1.5	90	60	<sup>35</sup> S	Mag spect	<sup>28</sup> APALIKOV 85
< 3.0	90	5–50		Mag spect	<sup>29</sup> MARKEY 85
< 0.62	90	48	<sup>35</sup> S	Si(Li)	<sup>30</sup> OHI 85
< 0.90	90	30	<sup>35</sup> S	Si(Li)	<sup>31</sup> OHI 85
< 4	90	140	<sup>64</sup> Cu	Mag spect	<sup>32</sup> SCHRECK... 83
< 8	90	440	<sup>64</sup> Cu	Mag spect	<sup>33</sup> SCHRECK... 83
<100	90	0.1–3000		THEO	<sup>34</sup> SHROCK 80
< 0.1	68	80		THEO	<sup>35</sup> SHROCK 80

• • • We do not use the following data for averages, fits, limits, etc. • • •

# Lepton Particle Listings

## Heavy Neutral Leptons, Searches for

15 TRINCEK 03 is a search for admixture of heavy neutrino to  $\nu_e$ , in contrast to  $\bar{\nu}_e$  used in many other searches. Full kinematic reconstruction of the neutrino momentum by use of a magneto optical trap.

16 GALEAZZI 01 use an cryogenic microcalorimeter to search for mass 50–1000 eV neutrino admixtures using the  $^{187}\text{Re}$  beta spectrum with 2.4 keV endpoint. They derive limits for the admixture of heavy neutrinos, ranging from  $9 \times 10^{-3}$  for mass 1 keV to 0.116 for mass 100 eV. This is a significant improvement with respect to HIDDEMANN 95, especially for masses below  $\sim 500$  MeV, where the limit is about a factor of  $\sim 2$  higher.

17 HOLZSCHUH 00 use an iron-free  $\beta$  spectrometer to measure the  $^{35}\text{S}$   $\beta$  decay spectrum. An analysis of the spectrum in the energy range 56–173 keV is used to derive limits for the admixture of heavy neutrinos. This extends the range of neutrino masses explored in HOLZSCHUH 99.

18 DRAGOUN 99 analyze the  $\beta$  decay spectrum of  $^{241}\text{Pu}$  in the energy range 0.2–9.2 keV to derive limits for the admixture of heavy neutrinos. It is not competitive with HOLZSCHUH 99.

19 HOLZSCHUH 99 use an iron-free  $\beta$  spectrometer to measure the  $^{63}\text{Ni}$   $\beta$  decay spectrum. An analysis of the spectrum in the energy range 33–67.8 keV is used to derive limits for the admixture of heavy neutrinos.

20 HINDI 98 obtain a limit on heavy neutrino admixture from EC decay of  $^{37}\text{Ar}$  by measuring the time-of-flight distribution of the recoiling ions in coincidence with x-rays or Auger electrons. The authors report upper limit for  $|U_{ex}|^2$  of  $\approx 3\%$  for  $m_{\nu_x}=500$  keV, 1% for  $m_{\nu_x}=550$  keV, 2% for  $m_{\nu_x}=600$  keV, and 4% for  $m_x=650$  keV. Their reported limits for  $m_{\nu_x} \leq 450$  keV are inferior to the limits of SCHRECKENBACH 83.

21 In the beta spectrum from tritium  $\beta$  decay nonvanishing or mixed  $m_{\bar{\nu}_1}$  state in the mass region 0.01–4 keV. For  $m_{\nu_x} < 1$  keV, their upper limit on  $|U_{ex}|^2$  becomes less

22 KALBFLEISCH 93 extends the 17 keV neutrino search of BAHRAN 92, using an improved proportional chamber to which a small amount of  $^3\text{H}$  is added. Systematics are significantly reduced, allowing for an improved upper limit. The authors give a 99% confidence limit on  $|U_{ex}|^2$  as a function of  $m_{\nu_x}$  in the range from 13.5 keV to 17.5 keV. See also the related papers BAHRAN 93, BAHRAN 93b, and BAHRAN 95 on theoretical aspects of beta spectra and fitting methods for heavy neutrinos.

23 MORTARA 93 limit is from study using a high-resolution solid-state detector with a superconducting solenoid. The authors note that “The sensitivity to neutrino mass is verified by measurement with a mixed source of  $^{35}\text{S}$  and  $^{14}\text{C}$ , which artificially produces a distortion in the beta spectrum similar to that expected from the massive neutrino.”

24 DEUTSCH 90 search for emission of heavy  $\bar{\nu}_e$  in super-allowed beta decay of  $^{20}\text{F}$  by spectral analysis of the electrons.

25 This limit was taken from the figure 3 of APALIKOV 85; the text gives a more restrictive limit of  $1.7 \times 10^{-3}$  at CL = 90%.

26 SCHRECKENBACH 83 is a combined measurement of the  $\beta^+$  and  $\beta^-$  spectrum.

27 SHROCK 80 was a retroactive analysis of data on several superallowed  $\beta$  decays to search for kinks in the Kurie plot.

28 Application of test to search for kinks in  $\beta$  decay Kurie plots.

### Searches for Decays of Massive $\nu$

Limits on  $|U_{ex}|^2$  as function of  $m_{\nu_x}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.6 \times 10^{-4}$	90	29 BACK	03A CNTR	$m_{\nu_x} = 4$ MeV
$<4.5 \times 10^{-5}$	90	29 BACK	03A CNTR	$m_{\nu_x} = 7$ MeV
$<3.8 \times 10^{-5}$	90	29 BACK	03A CNTR	$m_{\nu_x} = 10$ MeV
$<1.5 \times 10^{-3}$	95	ACHARD	01 L3	$m_{\nu_x}=80$ GeV
$<2 \times 10^{-2}$	95	ACHARD	01 L3	$m_{\nu_x}=175$ GeV
$<0.3$	95	ACHARD	01 L3	$m_{\nu_x}=200$ GeV
$<4 \times 10^{-3}$	95	ACCIARRI	99K L3	$m_{\nu_x}=80$ GeV
$<5 \times 10^{-2}$	95	ACCIARRI	99K L3	$m_{\nu_x}=175$ GeV
$<2 \times 10^{-5}$	95	30 ABREU	97I DLPH	$m_{\nu_x}=6$ GeV
$<3 \times 10^{-5}$	95	30 ABREU	97I DLPH	$m_{\nu_x}=50$ GeV
$<1.8 \times 10^{-3}$	90	31 HAGNER	95 MWPC	$m_{\nu_h} = 1.5$ MeV
$<2.5 \times 10^{-4}$	90	31 HAGNER	95 MWPC	$m_{\nu_h} = 4$ MeV
$<4.2 \times 10^{-3}$	90	31 HAGNER	95 MWPC	$m_{\nu_h} = 9$ MeV
$<1 \times 10^{-5}$	90	32 BARANOV	93	$m_{\nu_x}=100$ MeV
$<1 \times 10^{-6}$	90	32 BARANOV	93	$m_{\nu_x}=200$ MeV
$<3 \times 10^{-7}$	90	32 BARANOV	93	$m_{\nu_x}=300$ MeV
$<2 \times 10^{-7}$	90	32 BARANOV	93	$m_{\nu_x}=400$ MeV
$<6.2 \times 10^{-8}$	95	ADEVA	90s L3	$m_{\nu_x}=20$ GeV
$<5.1 \times 10^{-10}$	95	ADEVA	90s L3	$m_{\nu_x}=40$ GeV
all values ruled out	95	33 BURCHAT	90 MRK2	$m_{\nu_x} < 19.6$ GeV
$<1 \times 10^{-10}$	95	33 BURCHAT	90 MRK2	$m_{\nu_x} = 22$ GeV
$<1 \times 10^{-11}$	95	33 BURCHAT	90 MRK2	$m_{\nu_x} = 41$ GeV
all values ruled out	95	DECAMP	90f ALEP	$m_{\nu_x} = 25.0\text{--}42.7$ GeV
$<1 \times 10^{-13}$	95	DECAMP	90f ALEP	$m_{\nu_x} = 42.7\text{--}45.7$ GeV
$<5 \times 10^{-3}$	90	AKERLOF	88 HRS	$m_{\nu_x}=1.8$ GeV
$<2 \times 10^{-5}$	90	AKERLOF	88 HRS	$m_{\nu_x}=4$ GeV
$<3 \times 10^{-6}$	90	AKERLOF	88 HRS	$m_{\nu_x}=6$ GeV
$<1.2 \times 10^{-7}$	90	BERNARDI	88 CNTR	$m_{\nu_x}=100$ MeV

$<1 \times 10^{-8}$  90 BERNARDI 88 CNTR  $m_{\nu_x}=200$  MeV

$<2.4 \times 10^{-9}$  90 BERNARDI 88 CNTR  $m_{\nu_x}=300$  MeV

$<2.1 \times 10^{-9}$  90 BERNARDI 88 CNTR  $m_{\nu_x}=400$  MeV

$<2 \times 10^{-2}$  68 34 OBERAUER 87  $m_{\nu_x}=1.5$  MeV

$<8 \times 10^{-4}$  68 34 OBERAUER 87  $m_{\nu_x}=4.0$  MeV

$<8 \times 10^{-3}$  90 BADIER 86 CNTR  $m_{\nu_x}=400$  MeV

$<8 \times 10^{-5}$  90 BADIER 86 CNTR  $m_{\nu_x}=1.7$  GeV

$<8 \times 10^{-8}$  90 BERNARDI 86 CNTR  $m_{\nu_x}=100$  MeV

$<4 \times 10^{-8}$  90 BERNARDI 86 CNTR  $m_{\nu_x}=200$  MeV

$<6 \times 10^{-9}$  90 BERNARDI 86 CNTR  $m_{\nu_x}=400$  MeV

$<3 \times 10^{-5}$  90 DORENBOS... 86 CNTR  $m_{\nu_x}=150$  MeV

$<1 \times 10^{-6}$  90 DORENBOS... 86 CNTR  $m_{\nu_x}=500$  MeV

$<1 \times 10^{-7}$  90 DORENBOS... 86 CNTR  $m_{\nu_x}=1.6$  GeV

$<7 \times 10^{-7}$  90 35 COOPER... 85 HLBC  $m_{\nu_x}=0.4$  GeV

$<8 \times 10^{-8}$  90 35 COOPER... 85 HLBC  $m_{\nu_x}=1.5$  GeV

$<1 \times 10^{-2}$  90 36 BERGSMA 83B CNTR  $m_{\nu_x}=10$  MeV

$<1 \times 10^{-5}$  90 36 BERGSMA 83B CNTR  $m_{\nu_x}=110$  MeV

$<6 \times 10^{-7}$  90 36 BERGSMA 83B CNTR  $m_{\nu_x}=410$  MeV

$<1 \times 10^{-5}$  90 GRONAU 83  $m_{\nu_x}=160$  MeV

$<1 \times 10^{-6}$  90 GRONAU 83  $m_{\nu_x}=480$  MeV

29 BACK 03A searched for heavy neutrinos emitted from  $^8\text{B}$  decay in the Sun using the decay  $\nu_h \rightarrow \nu_e e^+ e^-$  in the Counting Test Facility (the prototype of the Borexino detector) and obtained limits on heavy neutrino admixture for the  $\nu_h$  mass range 1.1–12 MeV.

30 ABREU 97I long-lived  $\nu_x$  analysis. Short-lived analysis extends limit to lower masses with decreasing sensitivity except at 3.5 GeV, where the limit is the same as at 6 GeV.

31 HAGNER 95 obtain limits on heavy neutrino admixture from the decay  $\nu_h \rightarrow \nu_e e^+ e^-$  at a nuclear reactor for the  $\nu_h$  mass range 2–9 MeV.

32 BARANOV 93 is a search for neutrino decays into  $e^+ e^- \nu_e$  using a beam dump experiment at the 70 GeV Serpukhov proton synchrotron. The limits are not as good as those achieved earlier by BERGSMA 83 and BERNARDI 86, BERNARDI 88.

33 BURCHAT 90 includes the analyses reported in JUNG 90, ABRAMS 89c, and WENDT 87.

34 OBERAUER 87 bounds from search for  $\nu \rightarrow \nu' ee$  decay mode using reactor (anti)neutrinos.

35 COOPER-SARKAR 85 also give limits based on model-dependent assumptions for  $\nu_\tau$  flux. We do not list these. Note that for this bound to be nontrivial,  $x$  is not equal to 3, i.e.  $\nu_x$  cannot be the dominant mass eigenstate in  $\nu_\tau$  since  $m_{\nu_3} < 70$  MeV (ALBRECHT 85i). Also, of course,  $x$  is not equal to 1 or 2, so a fourth generation would be required for this bound to be nontrivial.

36 BERGSMA 83B also quote limits on  $|U_{e3}|^2$  where the index 3 refers to the mass eigenstate dominantly coupled to the  $\tau$ . Those limits were based on assumptions about the  $D_s$  mass and  $D_s \rightarrow \tau \nu_\tau$  branching ratio which are no longer valid. See COOPER-SARKAR 85.

### Limits on Coupling of $\mu$ to $\nu_x$ as Function of $m_{\nu_x}$

#### Peak search test

Limits on  $B(\pi(\text{ or } K) \rightarrow \mu \nu_x)$ .

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		37 ASTIER	02 NOMD	$\pi \rightarrow \mu X$ for $m_X=33.9$ MeV
$<6.0 \times 10^{-10}$	95	38 DAUM	00 CNTR	$\pi \rightarrow \mu X$ for $m_X=33.9$ MeV
		39 FORMAGGIO	00 CNTR	$\pi \rightarrow \mu X$ for $m_X=33.9$ MeV
$<0.22$	90	40 ASSAMAGAN	98 SILI	$m_{\nu_x} = 0.53$ MeV
$<0.029$	90	40 ASSAMAGAN	98 SILI	$m_{\nu_x} = 0.75$ MeV
$<0.016$	90	40 ASSAMAGAN	98 SILI	$m_{\nu_x} = 1.0$ MeV
$<4\text{--}6 \times 10^{-5}$		41 BRYMAN	96 CNTR	$m_{\nu_x} = 30\text{--}33.91$ MeV
$\sim 1 \times 10^{-16}$		42 ARMBRUSTER	95 KARM	$m_{\nu_x} = 33.9$ MeV
$<4 \times 10^{-7}$	95	43 BILGER	95 LEPS	$m_{\nu_x} = 33.9$ MeV
$<7 \times 10^{-8}$	95	43 BILGER	95 LEPS	$m_{\nu_x} = 33.9$ MeV
$<2.6 \times 10^{-8}$	95	43 DAUM	95b TOF	$m_{\nu_x} = 33.9$ MeV
$<2 \times 10^{-2}$	90	DAUM	87	$m_{\nu_x}=1$ MeV
$<1 \times 10^{-3}$	90	DAUM	87	$m_{\nu_x}=2$ MeV
$<6 \times 10^{-5}$	90	DAUM	87	$3 \text{ MeV} < m_{\nu_x} < 19.5$ MeV
$<3 \times 10^{-2}$	90	44 MINEHART	84	$m_{\nu_x}=2$ MeV
$<1 \times 10^{-3}$	90	44 MINEHART	84	$m_{\nu_x}=4$ MeV
$<3 \times 10^{-4}$	90	44 MINEHART	84	$m_{\nu_x}=10$ GeV
$<5 \times 10^{-6}$	90	45 HAYANO	82	$m_{\nu_x}=330$ MeV
$<1 \times 10^{-4}$	90	45 HAYANO	82	$m_{\nu_x}=70$ MeV
$<9 \times 10^{-7}$	90	45 HAYANO	82	$m_{\nu_x}=250$ MeV

See key on page 601

## Lepton Particle Listings

### Heavy Neutral Leptons, Searches for

<1	$\times 10^{-1}$	90	44 ABELA	81	$m_{\nu_X}=4$ MeV
<7	$\times 10^{-5}$	90	44 ABELA	81	$m_{\nu_X}=10.5$ MeV
<2	$\times 10^{-4}$	90	44 ABELA	81	$m_{\nu_X}=11.5$ MeV
<2	$\times 10^{-5}$	90	44 ABELA	81	$m_{\nu_X}=16-30$ MeV

<sup>37</sup> ASTIER 02 search for anomalous pion decay into a 33.9 MeV neutral particle. No evidence was found and the sensitivity to the branching ratio  $B(\pi \rightarrow \mu X) \cdot B(X \rightarrow \nu e^+ e^-)$  is as low as  $3.7 \times 10^{-15}$ , depending on the  $X$  lifetime.

<sup>38</sup> DAUM 00 search for anomalous pion decay into a 33.9 MeV neutral particle that might be responsible for the time-distribution anomaly observed by the KARMEN Collaboration.

<sup>39</sup> FORMAGGIO 00 search for anomalous pion decay into a 33.9 MeV neutral particle  $Q^0$  that might be responsible for the time-distribution anomaly observed by the KARMEN Collaboration. In the E815 (NuTeV) experiment at Fermilab no evidence was found, with sensitivity for the pion branching ratio  $B(\pi \rightarrow \mu Q^0) \cdot B(Q^0 \rightarrow \text{visible})$  as low as  $10^{-13}$ .

<sup>40</sup> ASSAMAGAN 98 obtain a limit on heavy neutrino admixture from  $\pi^+$  decay essentially at rest, by measuring with good resolution the momentum distribution of the muons. However, the search uses an ad hoc shape correction. The authors report upper limit for  $|U_{\mu X}|^2$  of 0.22 for  $m_{\nu} = 0.53$  MeV, 0.029 for  $m_{\nu} = 0.75$  MeV, and 0.016 for  $m_{\nu} = 1.0$  MeV at 90%CL.

<sup>41</sup> BRYMAN 96 search for massive unconventional neutrinos of mass  $m_{\nu_X}$  in  $\pi^+$  decay.

<sup>42</sup> ARMBRUSTER 95 study the reactions  $^{12}\text{C}(\nu_e, e^-)^{12}\text{N}$  and  $^{12}\text{C}(\nu, \nu')^{12}\text{C}^*$  induced by neutrinos from  $\pi^+$  and  $\mu^+$  decay at the ISIS neutron spallation source at the Rutherford-Appleton laboratory. An anomaly in the time distribution can be interpreted as the decay  $\pi^+ \rightarrow \mu^+ \nu_X$ , where  $\nu_X$  is a neutral weakly interacting particle with mass  $\approx 33.9$  MeV and spin 1/2. The lower limit to the branching ratio is a function of the lifetime of the new massive neutral particle, and reaches a minimum of a few  $\times 10^{-16}$  for  $\tau_X \sim 5$  s.

<sup>43</sup> From experiments of  $\pi^+$  and  $\pi^-$  decay in flight at PSI, to check the claim of the KARMEN Collaboration quoted above (ARMBRUSTER 95).

<sup>44</sup>  $\pi^+ \rightarrow \mu^+ \nu_\mu$  peak search experiment.

<sup>45</sup>  $K^+ \rightarrow \mu^+ \nu_\mu$  peak search experiment.

#### Peak search test

Limits on  $|U_{\mu X}|^2$  as function of  $m_{\nu_X}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1–10 $\times 10^{-4}$		46 BRYMAN	96	CNTR $m_{\nu_X} = 30-33.91$ MeV
<2 $\times 10^{-5}$	95	47 ASANO	81	$m_{\nu_X}=70$ MeV
<3 $\times 10^{-6}$	95	47 ASANO	81	$m_{\nu_X}=210$ MeV
<3 $\times 10^{-6}$	95	47 ASANO	81	$m_{\nu_X}=230$ MeV
<6 $\times 10^{-6}$	95	48 ASANO	81	$m_{\nu_X}=240$ MeV
<5 $\times 10^{-7}$	95	48 ASANO	81	$m_{\nu_X}=280$ MeV
<6 $\times 10^{-6}$	95	48 ASANO	81	$m_{\nu_X}=300$ MeV
<1 $\times 10^{-2}$	95	CALAPRICE	81	$m_{\nu_X}=7$ MeV
<3 $\times 10^{-3}$	95	49 CALAPRICE	81	$m_{\nu_X}=33$ MeV
<1 $\times 10^{-4}$	68	50 SHROCK	81	THEO $m_{\nu_X}=13$ MeV
<3 $\times 10^{-5}$	68	50 SHROCK	81	THEO $m_{\nu_X}=33$ MeV
<6 $\times 10^{-3}$	68	51 SHROCK	81	THEO $m_{\nu_X}=80$ MeV
<5 $\times 10^{-3}$	68	51 SHROCK	81	THEO $m_{\nu_X}=120$ MeV

<sup>46</sup> BRYMAN 96 search for massive unconventional neutrinos of mass  $m_{\nu_X}$  in  $\pi^+$  decay.

They interpret the result as an upper limit for the admixture of a heavy sterile or otherwise

<sup>47</sup>  $K^+ \rightarrow \mu^+ \nu_\mu$  peak search experiment.

<sup>48</sup> Analysis of experiment on  $K^+ \rightarrow \mu^+ \nu_\mu \nu_X \bar{\nu}_X$  decay.

<sup>49</sup>  $\pi^+ \rightarrow \mu^+ \nu_\mu$  peak search experiment.

<sup>50</sup> Analysis of magnetic spectrometer experiment, bubble chamber experiment, and emulsion experiment on  $\pi^+ \rightarrow \mu^+ \nu_\mu$  decay.

<sup>51</sup> Analysis of magnetic spectrometer experiment on  $K \rightarrow \mu, \nu_\mu$  decay.

#### Peak Search in Muon Capture

Limits on  $|U_{\mu X}|^2$  as function of  $m_{\nu_X}$

VALUE	CL%	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<1 $\times 10^{-1}$		DEUTSCH	83 $m_{\nu_X}=45$ MeV
<7 $\times 10^{-3}$		DEUTSCH	83 $m_{\nu_X}=70$ MeV
<1 $\times 10^{-1}$		DEUTSCH	83 $m_{\nu_X}=85$ MeV

#### Searches for Decays of Massive $\nu$

Limits on  $|U_{\mu X}|^2$  as function of  $m_{\nu_X}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

<5 $\times 10^{-7}$	90	52 VAITAITIS	99	CCFR $m_{\nu_X}=0.28$ GeV
<8 $\times 10^{-8}$	90	52 VAITAITIS	99	CCFR $m_{\nu_X}=0.37$ GeV
<5 $\times 10^{-7}$	90	52 VAITAITIS	99	CCFR $m_{\nu_X}=0.50$ GeV
<6 $\times 10^{-8}$	90	52 VAITAITIS	99	CCFR $m_{\nu_X}=1.50$ GeV
<2 $\times 10^{-5}$	95	53 ABREU	97i	DLPH $m_{\nu_X}=6$ GeV
<3 $\times 10^{-5}$	95	53 ABREU	97i	DLPH $m_{\nu_X}=50$ GeV
<3 $\times 10^{-6}$	90	GALLAS	95	CNTR $m_{\nu_X}=1$ GeV
<3 $\times 10^{-5}$	90	54 VILAIN	95c	CHM2 $m_{\nu_X}=2$ GeV
<6.2 $\times 10^{-8}$	95	ADEVA	90s	L3 $m_{\nu_X}=20$ GeV
<5.1 $\times 10^{-10}$	95	ADEVA	90s	L3 $m_{\nu_X}=40$ GeV
all values ruled out	95	55 BURCHAT	90	MRK2 $m_{\nu_X} < 19.6$ GeV
<1 $\times 10^{-10}$	95	55 BURCHAT	90	MRK2 $m_{\nu_X}=22$ GeV
<1 $\times 10^{-11}$	95	55 BURCHAT	90	MRK2 $m_{\nu_X}=41$ GeV
all values ruled out	95	DECAMP	90f	ALEP $m_{\nu_X}=25.0-42.7$ GeV
<1 $\times 10^{-13}$	95	DECAMP	90f	ALEP $m_{\nu_X}=42.7-45.7$ GeV
<5 $\times 10^{-3}$	90	AKERLOF	88	HRS $m_{\nu_X}=1.8$ GeV
<2 $\times 10^{-5}$	90	AKERLOF	88	HRS $m_{\nu_X}=4$ GeV
<3 $\times 10^{-6}$	90	AKERLOF	88	HRS $m_{\nu_X}=6$ GeV
<1 $\times 10^{-7}$	90	BERNARDI	88	CNTR $m_{\nu_X}=200$ MeV
<3 $\times 10^{-9}$	90	BERNARDI	88	CNTR $m_{\nu_X}=300$ MeV
<4 $\times 10^{-4}$	90	56 MISHRA	87	CNTR $m_{\nu_X}=1.5$ GeV
<4 $\times 10^{-3}$	90	56 MISHRA	87	CNTR $m_{\nu_X}=2.5$ GeV
<0.9 $\times 10^{-2}$	90	56 MISHRA	87	CNTR $m_{\nu_X}=5$ GeV
<0.1	90	56 MISHRA	87	CNTR $m_{\nu_X}=10$ GeV
<8 $\times 10^{-4}$	90	BADIER	86	CNTR $m_{\nu_X}=600$ MeV
<1.2 $\times 10^{-5}$	90	BADIER	86	CNTR $m_{\nu_X}=1.7$ GeV
<3 $\times 10^{-8}$	90	BERNARDI	86	CNTR $m_{\nu_X}=200$ MeV
<6 $\times 10^{-9}$	90	BERNARDI	86	CNTR $m_{\nu_X}=350$ MeV
<1 $\times 10^{-6}$	90	DORENBOS...	86	CNTR $m_{\nu_X}=500$ MeV
<1 $\times 10^{-7}$	90	DORENBOS...	86	CNTR $m_{\nu_X}=1600$ MeV
<0.8 $\times 10^{-5}$	90	57 COOPER...	85	HLBC $m_{\nu_X}=0.4$ GeV
<1.0 $\times 10^{-7}$	90	57 COOPER...	85	HLBC $m_{\nu_X}=1.5$ GeV

<sup>52</sup> VAITAITIS 99 search for  $L_\mu^0 \rightarrow \mu X$ . See paper for rather complicated limit as function of  $m_{\nu_X}$ .

<sup>53</sup> ABREU 97i long-lived  $\nu_X$  analysis. Short-lived analysis extends limit to lower masses with decreasing sensitivity except at 3.5 GeV, where the limit is the same as at 6 GeV.

<sup>54</sup> VILAIN 95c is a search for the decays of heavy isosinglet neutrinos produced by neutral current neutrino interactions. Limits were quoted for masses in the range from 0.3 to 24 GeV. The best limit is listed above.

<sup>55</sup> BURCHAT 90 includes the analyses reported in JUNG 90, ABRAMS 89c, and WENDT 87.

<sup>56</sup> See also limits on  $|U_{3x}|$  from WENDT 87.

<sup>57</sup> COOPER-SARKAR 85 also give limits based on model-dependent assumptions for  $\nu_\tau$  flux. We do not list these. Note that for this bound to be nontrivial,  $x$  is not equal to 3, i.e.  $\nu_X$  cannot be the dominant mass eigenstate in  $\nu_\tau$  since  $m_{\nu_3} < 70$  MeV (ALBRECHT 85i). Also, of course,  $x$  is not equal to 1 or 2, so a fourth generation would be required for this bound to be nontrivial.

#### Limits on $|U_{\tau X}|^2$ as a Function of $m_{\nu_X}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1 $\times 10^{-2}$	90	58 ORLOFF	02	CHRM $m_{\nu_X}=45$ MeV
<1.4 $\times 10^{-4}$	90	58 ORLOFF	02	CHRM $m_{\nu_X}=180$ MeV
<0.025	90	ASTIER	01	$m_{\nu_X}=45$ MeV
<0.002	90	ASTIER	01	$m_{\nu_X}=140$ MeV
<2 $\times 10^{-5}$	95	59 ABREU	97i	DLPH $m_{\nu_X}=6$ GeV
<3 $\times 10^{-5}$	95	59 ABREU	97i	DLPH $m_{\nu_X}=50$ GeV
<6.2 $\times 10^{-8}$	95	ADEVA	90s	L3 $m_{\nu_X}=20$ GeV
<5.1 $\times 10^{-10}$	95	ADEVA	90s	L3 $m_{\nu_X}=40$ GeV
all values ruled out	95	60 BURCHAT	90	MRK2 $m_{\nu_X} < 19.6$ GeV
<1 $\times 10^{-10}$	95	60 BURCHAT	90	MRK2 $m_{\nu_X}=22$ GeV
<1 $\times 10^{-11}$	95	60 BURCHAT	90	MRK2 $m_{\nu_X}=41$ GeV
all values ruled out	95	DECAMP	90f	ALEP $m_{\nu_X}=25.0-42.7$ GeV
<1 $\times 10^{-13}$	95	DECAMP	90f	ALEP $m_{\nu_X}=42.7-45.7$ GeV
<5 $\times 10^{-2}$	80	AKERLOF	88	HRS $m_{\nu_X}=2.5$ GeV
<9 $\times 10^{-5}$	80	AKERLOF	88	HRS $m_{\nu_X}=4.5$ GeV

<sup>58</sup> ORLOFF 02 use the negative result of a search for neutral particles decaying into two electrons performed by CHARM to get these limits for a mostly isosinglet heavy neutrino.

<sup>59</sup> ABREU 97i long-lived  $\nu_X$  analysis. Short-lived analysis extends limit to lower masses with decreasing sensitivity.

<sup>60</sup> BURCHAT 90 includes the analyses reported in JUNG 90, ABRAMS 89c, and WENDT 87.

# Lepton Particle Listings

## Heavy Neutral Leptons, Searches for

### Limits on $|U_{\nu x}|^2$

Where  $a = e, \mu$  from  $\rho$  parameter in  $\mu$  decay.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1 \times 10^{-2}$	68	SHROCK	81B	THEO $m_{\nu_x}=10$ GeV
$<2 \times 10^{-3}$	68	SHROCK	81B	THEO $m_{\nu_x}=40$ MeV
$<4 \times 10^{-2}$	68	SHROCK	81B	THEO $m_{\nu_x}=70$ MeV

### Limits on $|U_{1j} \times U_{2j}|$ as Function of $m_{\nu_j}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3 \times 10^{-5}$	90	<sup>61</sup> BARANOV	93	$m_{\nu_j}=80$ MeV
$<3 \times 10^{-6}$	90	<sup>61</sup> BARANOV	93	$m_{\nu_j}=160$ MeV
$<6 \times 10^{-7}$	90	<sup>61</sup> BARANOV	93	$m_{\nu_j}=240$ MeV
$<2 \times 10^{-7}$	90	<sup>61</sup> BARANOV	93	$m_{\nu_j}=320$ MeV
$<9 \times 10^{-5}$	90	BERNARDI	86	CNTR $m_{\nu_j}=25$ MeV
$<3.6 \times 10^{-7}$	90	BERNARDI	86	CNTR $m_{\nu_j}=100$ MeV
$<3 \times 10^{-8}$	90	BERNARDI	86	CNTR $m_{\nu_j}=200$ MeV
$<6 \times 10^{-9}$	90	BERNARDI	86	CNTR $m_{\nu_j}=350$ MeV
$<1 \times 10^{-2}$	90	BERGSMA	83B	CNTR $m_{\nu_j}=10$ MeV
$<1 \times 10^{-5}$	90	BERGSMA	83B	CNTR $m_{\nu_j}=140$ MeV
$<7 \times 10^{-7}$	90	BERGSMA	83B	CNTR $m_{\nu_j}=370$ MeV

<sup>61</sup>BARANOV 93 is a search for neutrino decays into  $e^+ e^- \nu_e$  using a beam dump experiment at the 70 GeV Serpukhov proton synchrotron.

### REFERENCES FOR Heavy Neutral Leptons, Searches for

BACK	03A	JETPL 78 261	H.O. Back <i>et al.</i>	(Borexino Collab.)
TRINCZEK	03	Translated from ZETFP 78 707	M. Trinczek <i>et al.</i>	
ASTIER	02	PL B527 23	P. Astier <i>et al.</i>	(NOMAD Collab.)
ORLOFF	02	PL B550 8	J. Orloff <i>et al.</i>	
ACHARD	01	PL B517 67	P. Achard <i>et al.</i>	(L3 Collab.)
ACHARD	01B	PL B517 75	P. Achard <i>et al.</i>	(L3 Collab.)
ASTIER	01	PL B506 27	P. Astier <i>et al.</i>	(NOMAD Collab.)
GALEAZZI	01	PRL 86 1978	M. Galeazzi <i>et al.</i>	
ABBIENDI	00I	EPJ C14 73	G. Abbiendi <i>et al.</i>	
DAUM	00	PRL 85 1815	M. Daum <i>et al.</i>	(OPAL Collab.)
FORMAGGIO	00	PRL 84 4043	J.A. Formaggio <i>et al.</i>	
HOLZSCHUH	00	PL B482 1	E. Holzschuh <i>et al.</i>	
ABREU	99O	EPJ C8 41	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	99K	PL B461 397	M. Acciarri <i>et al.</i>	(L3 Collab.)
DRAGON	99	JP G25 1839	O. Dragoun <i>et al.</i>	
HOLZSCHUH	99	PL B451 247	E. Holzschuh <i>et al.</i>	
VAITAITIS	99	PRL 83 4943	A. Vaitaitis <i>et al.</i>	(CCFR Collab.)
ASSAMAGAN	98	PL B434 158	K. Assamagan <i>et al.</i>	
HINDI	98	PR C58 2512	M.M. Hindi <i>et al.</i>	
PDG	98	EPJ C3 1	C. Caso <i>et al.</i>	
ABREU	97I	ZPHY C74 57	P. Abreu <i>et al.</i>	(PDG Collab.)
Also		ZPHY C75 580 (erratum)	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BRYMAN	96	PR D53 558	D.A. Bryman, T. Numao	(DELPHI Collab.)
BUSKULIC	96S	PL B384 439	D. Buskulić <i>et al.</i>	(TRIUMF Collab.)
WIETFELDT	96	PRPL 273 149	F.E. Wietfeldt, E.B. Norman	(ALEPH Collab.)

ARMBRUSTER	95	PL B348 19	B. Armbruster <i>et al.</i>	(KARMEN Collab.)
BAHRAN	95	PL B354 481	M.Y. Bahrán, G.R. Kalbfleisch	(OKLA)
BILGER	95	PL B363 41	R. Bilger <i>et al.</i>	(TUBIN, KARLE, PSI)
DAUM	95B	PL B361 179	M. Daum <i>et al.</i>	(PSI, UVA)
FARGION	95	PR D52 1828	D. Fargion <i>et al.</i>	(ROMA, KIAM, MPEI)
GALLAS	95	PR D52 6	E. Gallas <i>et al.</i>	(MSU, FNAL, MIT, FLOR)
GARCIA	95	PR D51 1458	E. Garcia <i>et al.</i>	(ZARA, SCUC, PNL)
HAGNER	95	PR D52 1343	C. Hagner <i>et al.</i>	(MUNT, LAPP, CERN)
HIDDEMANN	95	JP G21 639	K.H. Hidde mann, H. Daniel, O. Schwenker	(MUNT)
VILAIN	95C	PL B351 387	P. Vilain <i>et al.</i>	(CHARM II Collab.)
Also		PL B343 453	P. Vilain <i>et al.</i>	(CHARM II Collab.)
BECK	94	PL B336 141	M. Beck <i>et al.</i>	(MPIH, KIAE, SASSO)
KONOPLICH	94	PAN 57 425	R.V. Konoplich, M.Y. Khlopov	(MPEI)
PDG	94	PR D50 1173	L. Montanet <i>et al.</i>	(CERN, LBL, BOST+)
BAHRAN	93	PR D47 R754	M. Bahrán, G.R. Kalbfleisch	(OKLA)
BAHRAN	93B	PR D47 R759	M. Bahrán, G.R. Kalbfleisch	(OKLA)
BARANOV	93	PL B302 336	S.A. Baranov <i>et al.</i>	(JINR, SERP, BUDA)
KALBFLEISCH	93	PL B303 355	G.R. Kalbfleisch, M.Y. Bahrán	(OKLA)
MORTARA	93	PRL 70 394	J.L. Mortara <i>et al.</i>	(ANL, LBL, UCB)
OSHIMA	93	PR D47 4840	T. Oshima <i>et al.</i>	(KEK, TUAT, RIKEN+)
ABREU	92B	PL B274 230	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BAHRAN	92	PL B291 336	M.Y. Bahrán, G.R. Kalbfleisch	(OKLA)
BRITTON	92	PRL 68 3000	D.I. Britton <i>et al.</i>	(TRIUMF, CARL)
Also		PR D49 28	D.I. Britton <i>et al.</i>	(TRIUMF, CARL)
BRITTON	92B	PR D46 R885	D.I. Britton <i>et al.</i>	(TRIUMF, CARL)
KAWAKAMI	92	PL B287 45	H. Kawakami <i>et al.</i>	(INUS, KEK, SCUC+)
MORI	92B	PL B289 463	M. Mori <i>et al.</i>	(KAM2 Collab.)
ALEXANDER	91F	ZPHY C52 175	G. Alexander <i>et al.</i>	(OPAL Collab.)
DELEENER...	91	PR D43 3611	N. de Leener-Rosier <i>et al.</i>	(LOUV, ZURICH)
REUSSER	91	PL B255 143	D. Reusser <i>et al.</i>	(NEUC, CIT, PSI)
SATO	91	PR D44 2220	N. Sato <i>et al.</i>	(Kamiokande Collab.)
ADEVA	90S	PL B251 321	B. Adeva <i>et al.</i>	(L3 Collab.)
BURCHAT	90	PR D41 3542	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
DECAMP	90F	PL B236 511	D. Decamp <i>et al.</i>	(ALEPH Collab.)
DEUTSCH	90	NP A518 149	J. Deutsch, M. Lebrun, R. Prieels	
JUNG	90	PRL 64 1091	C. Jung <i>et al.</i>	(Mark II Collab.)
ABRAMS	89C	PRL 63 2447	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
ENQVIST	89	NP B317 647	K. Enqvist, K. Kainulainen, J. Maalampi	(HELS)
FISHER	89	PL B218 257	P.H. Fisher <i>et al.</i>	(CIT, NEUC, PSI)
AKERLOF	88	PR D37 577	C.W. Akerlof <i>et al.</i>	(HRS Collab.)
BERNARDI	88	PL B203 332	G. Bernardi <i>et al.</i>	(PARIN, CERN, INFN+)
CALDWELL	88	PRL 61 510	D.O. Caldwell <i>et al.</i>	(UCSB, UCB, LBL)
OLIVE	88	PL B205 553	K.A. Olive, M. Srednicki	(MINN, UCSB)
SREDNICKI	88	NP B310 693	M. Srednicki, R. Watkins, K.A. Olive	(MINN, UCSB)
AHLEN	87	PL B195 603	S.P. Ahlen <i>et al.</i>	(BOST, SCUC, HARV+)
DAUM	87	PR D36 2624	M. Daum <i>et al.</i>	(SIN, UVA)
GRIEST	87	NP B283 681	K. Griest, D. Seckel	(UCSC, CERN)
Also		NP B296 1034 (erratum)	K. Griest, D. Seckel	(UCSC, CERN)
MISHRA	87	PRL 59 1397	S.R. Mishra <i>et al.</i>	(COLU, CIT, FNAL+)
OBERAUER	87	PL B198 113	L.F. Oberauer, F. von Feilitzsch, R.L. Mossbauer	
WENDT	87	PRL 58 1810	C. Wendt <i>et al.</i>	(Mark II Collab.)
AZUELOS	86	PRL 56 2241	G. Azuelos <i>et al.</i>	(TRIUMF, CNRC)
BADIER	86	ZPHY C51 21	J. Badier <i>et al.</i>	(NAS Collab.)
BERNARDI	86	PL B166 479	G. Bernardi <i>et al.</i>	(CURIN, INFN, CDF+)
DORNBOSCH...	86	PL B166 473	J. Dornbosch <i>et al.</i>	(CHARM Collab.)
ALBRECHT	85I	PL B163B 404	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
APALIKOV	85	JETPL 42 289	A.M. Apalikov <i>et al.</i>	(ITEP)
COOPER...	85	Translated from ZETFP 42 233	A.M. Cooper-Sarkar <i>et al.</i>	(CERN, LOIC+)
MARKET	85	PL B160B 207	J. Markey, F. Boehm	(CIT)
OHI	85	PL B160B 322	T. Ohi <i>et al.</i>	(TOKY, INUS, KEK)
MINEHART	84	PRL 52 804	R.C. Minehart <i>et al.</i>	(UVA, SIN)
BERGSMA	83	PL B122B 465	F. Bergsma <i>et al.</i>	(CHARM Collab.)
BERGSMA	83B	PL B128B 361	F. Bergsma <i>et al.</i>	(CHARM Collab.)
BRYMAN	83B	PRL 50 1546	D.A. Bryman <i>et al.</i>	(TRIUMF, CNRC)
DEUTSCH	83	PR D27 1644	J.P. Deutsch, M. Lebrun, R. Prieels	(LOUV)
GRONAU	83	PR D28 2762	M. Gronau	(HAIF)
SCHRECK...	83	PL B129B 265	K. Schreckenbach <i>et al.</i>	(ISNG, ILG)
HAYANO	82	PRL 49 1305	R.S. Hayano <i>et al.</i>	(TOKY, KEK, TSUK)
ABELA	81	PL B105B 263	R. Abela <i>et al.</i>	(SIN)
ASANO	81	PL B104B 84	Y. Asano <i>et al.</i>	(KEK, TOKY, INUS, OSAK)
CALAPRICE	81	PL B106B 175	F.P. Calaprice <i>et al.</i>	(PRIN, IND)
SHROCK	81	PR D24 1232	R.E. Shrock	(STON)
SHROCK	81B	PR D24 1275	R.E. Shrock	(STON)
SHROCK	80	PL B6B 159	R.E. Shrock	(STON)

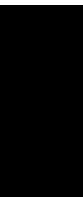
**QUARKS**

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<i>s</i> . . . . .	801
<i>c</i> . . . . .	804
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## QUARKS

**QUARK MASSES**

Updated Jan 2016 by A.V. Manohar (University of California, San Diego), C.T. Sachrajda (University of Southampton), and R.M. Barnett (LBNL).

**A. Introduction**

This note discusses some of the theoretical issues relevant for the determination of quark masses, which are fundamental parameters of the Standard Model of particle physics. Unlike the leptons, quarks are confined inside hadrons and are not observed as physical particles. Quark masses therefore cannot be measured directly, but must be determined indirectly through their influence on hadronic properties. Although one often speaks loosely of quark masses as one would of the mass of the electron or muon, any quantitative statement about the value of a quark mass must make careful reference to the particular theoretical framework that is used to define it. It is important to keep this *scheme dependence* in mind when using the quark mass values tabulated in the data listings.

Historically, the first determinations of quark masses were performed using quark models. The resulting masses only make sense in the limited context of a particular quark model, and cannot be related to the quark mass parameters of the Standard Model. In order to discuss quark masses at a fundamental level, definitions based on quantum field theory must be used, and the purpose of this note is to discuss these definitions and the corresponding determinations of the values of the masses.

**B. Mass parameters and the QCD Lagrangian**

The QCD [1] Lagrangian for  $N_F$  quark flavors is

$$\mathcal{L} = \sum_{k=1}^{N_F} \bar{q}_k (i \not{D} - m_k) q_k - \frac{1}{4} G_{\mu\nu} G^{\mu\nu}, \quad (1)$$

where  $\not{D} = (\partial_\mu - ig A_\mu) \gamma^\mu$  is the gauge covariant derivative,  $A_\mu$  is the gluon field,  $G_{\mu\nu}$  is the gluon field strength,  $m_k$  is the mass parameter of the  $k^{\text{th}}$  quark, and  $q_k$  is the quark Dirac field. After renormalization, the QCD Lagrangian Eq. (1) gives finite values for physical quantities, such as scattering amplitudes. Renormalization is a procedure that invokes a subtraction scheme to render the amplitudes finite, and requires the introduction of a dimensionful scale parameter  $\mu$ . The mass parameters in the QCD Lagrangian Eq. (1) depend on the renormalization scheme used to define the theory, and also on the scale parameter  $\mu$ . The most commonly used renormalization scheme for QCD perturbation theory is the  $\overline{\text{MS}}$  scheme.

The QCD Lagrangian has a chiral symmetry in the limit that the quark masses vanish. This symmetry is spontaneously

broken by dynamical chiral symmetry breaking, and explicitly broken by the quark masses. The nonperturbative scale of dynamical chiral symmetry breaking,  $\Lambda_\chi$ , is around 1 GeV [2]. It is conventional to call quarks heavy if  $m > \Lambda_\chi$ , so that explicit chiral symmetry breaking dominates ( $c$ ,  $b$ , and  $t$  quarks are heavy), and light if  $m < \Lambda_\chi$ , so that spontaneous chiral symmetry breaking dominates (the  $u$  and  $d$  are light and  $s$  quarks are considered to be light when using  $\text{SU}(3)_L \times \text{SU}(3)_R$  chiral perturbation theory). The determination of light- and heavy-quark masses is considered separately in sections D and E below.

At high energies or short distances, nonperturbative effects, such as chiral symmetry breaking, become small and one can, in principle, determine quark masses by analyzing mass-dependent effects using QCD perturbation theory. Such computations are conventionally performed using the  $\overline{\text{MS}}$  scheme at a scale  $\mu \gg \Lambda_\chi$ , and give the  $\overline{\text{MS}}$  “running” mass  $\overline{m}(\mu)$ . We use the  $\overline{\text{MS}}$  scheme when reporting quark masses; one can readily convert these values into other schemes using perturbation theory.

The  $\mu$  dependence of  $\overline{m}(\mu)$  at short distances can be calculated using the renormalization group equation,

$$\mu^2 \frac{d\overline{m}(\mu)}{d\mu^2} = -\gamma(\overline{\alpha}_s(\mu)) \overline{m}(\mu), \quad (2)$$

where  $\gamma$  is the anomalous dimension which is now known to four-loop order in perturbation theory [3,4].  $\overline{\alpha}_s$  is the coupling constant in the  $\overline{\text{MS}}$  scheme. Defining the expansion coefficients  $\gamma_r$  by

$$\gamma(\overline{\alpha}_s) \equiv \sum_{r=1}^{\infty} \gamma_r \left( \frac{\overline{\alpha}_s}{4\pi} \right)^r,$$

the first four coefficients are given by

$$\begin{aligned} \gamma_1 &= 4, \\ \gamma_2 &= \frac{202}{3} - \frac{20N_L}{9}, \\ \gamma_3 &= 1249 + \left( -\frac{2216}{27} - \frac{160}{3}\zeta(3) \right) N_L - \frac{140}{81} N_L^2, \\ \gamma_4 &= \frac{4603055}{162} + \frac{135680}{27}\zeta(3) - 8800\zeta(5) \\ &\quad + \left( -\frac{91723}{27} - \frac{34192}{9}\zeta(3) + 880\zeta(4) + \frac{18400}{9}\zeta(5) \right) N_L \\ &\quad + \left( \frac{5242}{243} + \frac{800}{9}\zeta(3) - \frac{160}{3}\zeta(4) \right) N_L^2 \\ &\quad + \left( -\frac{332}{243} + \frac{64}{27}\zeta(3) \right) N_L^3, \end{aligned}$$

where  $N_L$  is the number of active light quark flavors at the scale  $\mu$ , i.e. flavors with masses  $< \mu$ , and  $\zeta$  is the Riemann

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zeta function ( $\zeta(3) \simeq 1.2020569$ ,  $\zeta(4) \simeq 1.0823232$ , and  $\zeta(5) \simeq 1.0369278$ ). In addition, as the renormalization scale crosses quark mass thresholds one needs to match the scale dependence of  $\overline{m}$  below and above the threshold. There are finite threshold corrections; the necessary formulae can be found in Ref. [5].

The quark masses for light quarks discussed so far are often referred to as current quark masses. Nonrelativistic quark models use constituent quark masses, which are of order 350 MeV for the  $u$  and  $d$  quarks. Constituent quark masses model the effects of dynamical chiral symmetry breaking, and are not directly related to the quark mass parameters  $m_k$  of the QCD Lagrangian Eq. (1). Constituent masses are only defined in the context of a particular hadronic model.

### C. Lattice Gauge Theory

The use of the lattice simulations for *ab initio* determinations of the fundamental parameters of QCD, including the coupling constant and quark masses (except for the top-quark mass) is a very active area of research (see the review on Lattice Quantum Chromodynamics in this *Review*). Here we only briefly recall those features which are required for the determination of quark masses. In order to determine the lattice spacing ( $a$ , i.e. the distance between neighboring points of the lattice) and quark masses, one computes a convenient and appropriate set of physical quantities (frequently chosen to be a set of hadronic masses) for a variety of input values of the quark masses. The true (physical) values of the quark masses are those which correctly reproduce the set of physical quantities being used for the calibration.

The values of the quark masses obtained directly in lattice simulations are bare quark masses, corresponding to a particular discretization of QCD and with the lattice spacing as the ultraviolet cut-off. In order for these results to be useful in phenomenological applications, it is necessary to relate them to renormalized masses defined in some standard renormalization scheme such as  $\overline{\text{MS}}$ . Provided that both the ultraviolet cut-off  $a^{-1}$  and the renormalization scale  $\mu$  are much greater than  $\Lambda_{\text{QCD}}$ , the bare and renormalized masses can be related in perturbation theory. However, in order to avoid uncertainties due to the unknown higher-order coefficients in lattice perturbation theory, most results obtained recently use *non-perturbative renormalization* to relate the bare masses to those defined in renormalization schemes which can be simulated directly in lattice QCD (e.g. those obtained from quark and gluon Green functions at specified momenta in the Landau gauge [62] or those defined using finite-volume techniques and the Schrödinger functional [63]). The conversion to the  $\overline{\text{MS}}$  scheme (which cannot be simulated) is then performed using continuum perturbation theory.

The determination of quark masses using lattice simulations is well established and the current emphasis is on the reduction and control of the systematic uncertainties. With improved algorithms and access to more powerful computing resources, the precision of the results has improved immensely in recent years.

Vacuum polarisation effects are included with  $N_f = 2, 2 + 1$  or  $N_f = 2 + 1 + 1$  flavors of sea quarks. The number 2 here indicates that the up and down quarks are degenerate. In earlier *reviews*, results were presented from simulations in which vacuum polarization effects were completely neglected (this is the so-called *quenched* approximation), leading to systematic uncertainties which could not be estimated reliably. It is no longer necessary to include quenched results in compilations of quark masses. Particularly pleasing is the observation that results obtained using different formulations of lattice QCD, with different systematic uncertainties, give results which are largely consistent with each other. This gives us broad confidence in the estimates of the systematic errors. As the precision of the results approaches (or even exceeds in some cases) 1%, isospin breaking effects, including electromagnetic corrections need to be included and this is beginning to be done as will be discussed below. The results however, are still at an early stage and therefore, unless explicitly stated otherwise, the results presented below will neglect isospin breaking.

Members of the lattice QCD community have organised a Flavour Lattice Averaging Group (FLAG) which critically reviews quantities computed in lattice QCD relevant to flavor physics, including the determination of light quark masses, against stated quality criteria and presents its view of the current status of the results. The latest (2nd) edition reviewed lattice results published before November 30th 2013 [16].

### D. Light quarks

In this section we review the determination of the masses of the light quarks  $u$ ,  $d$  and  $s$  from lattice simulations and then discuss the consequences of the approximate chiral symmetry.

**Lattice Gauge Theory:** The most reliable determinations of the strange quark mass  $m_s$  and of the average of the up and down quark masses  $m_{ud} = (m_u + m_d)/2$  are obtained from lattice simulations. As explained in section C above, the simulations are generally performed with degenerate up and down quarks ( $m_u = m_d$ ) and so it is the average which is obtained directly from the computations. Below we discuss attempts to derive  $m_u$  and  $m_d$  separately using lattice results in combination with other techniques, but we start by briefly present our estimate of the current status of the latest lattice results in the isospin symmetric limit. Based largely on references [21–25], which its authors considered to have the most reliable estimates of the systematic uncertainties, the FLAG Review [16] quoted as its summary of results obtained with  $N_f = 2 + 1$  flavors of sea quarks:

$$\overline{m}_s = (93.8 \pm 1.5 \pm 1.9) \text{ MeV}, \quad (3)$$

$$\overline{m}_{ud} = (3.42 \pm 0.06 \pm 0.07) \text{ MeV} \quad (4)$$

and

$$\frac{\overline{m}_s}{\overline{m}_{ud}} = 27.46 \pm 0.15 \pm 0.41. \quad (5)$$

The masses are given in the  $\overline{\text{MS}}$  scheme at a renormalization scale of 2 GeV. The first error comes from averaging the lattice results and the second is an estimate of the neglect of sea-quark effects from the charm and more massive quarks. Because of the systematic errors, these results are not simply the combinations of all the results in quadrature, but include a judgement of the remaining uncertainties. Since the different collaborations use different formulations of lattice QCD, the (relatively small) variations of the results between the groups provides important information about the reliability of the estimates.

Since the publication of the FLAG review [16] there have been a number of studies with  $N_f = 2 + 1 + 1$  [26–28] and  $N_f = 2 + 1$  [29] and a reasonable summary of the current status may be  $\overline{m}_{ud} = (3.4 \pm 0.1) \text{ MeV}$ ,  $\overline{m}_s = (93.5 \pm 2) \text{ MeV}$  and  $\overline{m}_s/\overline{m}_{ud} = 27.5 \pm 0.3$ .

To obtain the individual values of  $\overline{m}_u$  and  $\overline{m}_d$  requires the introduction of isospin breaking effects, including electromagnetism. In principle this can be done completely using lattice field theory. Such calculations are indeed beginning (note the recent computation of the neutron-proton mass splitting [30]) but are still at a relatively early stage. In practice therefore,  $\overline{m}_u$  and  $\overline{m}_d$  are extracted by combining lattice results with some elements of continuum phenomenology, most frequently based on chiral perturbation theory. Such studies include references [32,17,24,28,33,34] as well the Flavianet Lattice Averaging Group [43]. Based on these results we summarise the current status as

$$\frac{\overline{m}_u}{\overline{m}_d} = 0.46(5), \quad \overline{m}_u = 2.15(15) \text{ MeV}, \quad \overline{m}_d = 4.70(20) \text{ MeV}. \quad (6)$$

Again the masses are given in the  $\overline{\text{MS}}$  scheme at a renormalization scale of 2 GeV. Of particular importance is the fact that  $m_u \neq 0$  since there would have been no strong  $CP$  problem had  $m_u$  been equal to zero.

The quark mass ranges for the light quarks given in the listings combine the lattice and continuum values and use the PDG method for determining errors given in the introductory notes.

**Chiral Perturbation Theory:** For light quarks, one can use the techniques of chiral perturbation theory [6–8] to extract quark mass ratios. The mass term for light quarks in the QCD Lagrangian is

$$\overline{\Psi} M \Psi = \overline{\Psi}_L M \Psi_R + \overline{\Psi}_R M^\dagger \Psi_L, \quad (7)$$

where  $M$  is the light quark mass matrix,

$$M = \begin{pmatrix} m_u & 0 & 0 \\ 0 & m_d & 0 \\ 0 & 0 & m_s \end{pmatrix}, \quad (8)$$

$\Psi = (u, d, s)$ , and  $L$  and  $R$  are the left- and right-chiral components of  $\Psi$  given by  $\Psi_{L,R} = P_{L,R} \Psi$ ,  $P_L = (1 - \gamma_5)/2$ ,  $P_R = (1 + \gamma_5)/2$ . The mass term is the only term in the QCD Lagrangian that mixes left- and right-handed quarks. In the

limit  $M \rightarrow 0$ , there is an independent  $SU(3) \times U(1)$  flavor symmetry for the left- and right-handed quarks. The vector  $U(1)$  symmetry is baryon number; the axial  $U(1)$  symmetry of the classical theory is broken in the quantum theory due to the anomaly. The remaining  $G_\chi = SU(3)_L \times SU(3)_R$  chiral symmetry of the QCD Lagrangian is spontaneously broken to  $SU(3)_V$ , which, in the limit  $M \rightarrow 0$ , leads to eight massless Goldstone bosons, the  $\pi$ 's,  $K$ 's, and  $\eta$ .

The symmetry  $G_\chi$  is only an approximate symmetry, since it is explicitly broken by the quark mass matrix  $M$ . The Goldstone bosons acquire masses which can be computed in a systematic expansion in  $M$ , in terms of low-energy constants, which are unknown nonperturbative parameters of the effective theory, and are not fixed by the symmetries. One treats the quark mass matrix  $M$  as an external field that transforms under  $G_\chi$  as  $M \rightarrow L M R^\dagger$ , where  $\Psi_L \rightarrow L \Psi_L$  and  $\Psi_R \rightarrow R \Psi_R$  are the  $SU(3)_L$  and  $SU(3)_R$  transformations, and writes down the most general Lagrangian invariant under  $G_\chi$ . Then one sets  $M$  to its given constant value Eq. (8), which implements the symmetry breaking. To first order in  $M$  one finds that [9]

$$\begin{aligned} m_{\pi^0}^2 &= B(m_u + m_d), \\ m_{\pi^\pm}^2 &= B(m_u + m_d) + \Delta_{\text{em}}, \\ m_{K^0}^2 &= m_{K^\pm}^2 = B(m_d + m_s), \\ m_{K^\pm}^2 &= B(m_u + m_s) + \Delta_{\text{em}}, \\ m_\eta^2 &= \frac{1}{3} B(m_u + m_d + 4m_s), \end{aligned} \quad (9)$$

with two unknown constants  $B$  and  $\Delta_{\text{em}}$ , the electromagnetic mass difference. From Eq. (9), one can determine the quark mass ratios [9]

$$\begin{aligned} \frac{m_u}{m_d} &= \frac{2m_{\pi^0}^2 - m_{\pi^+}^2 + m_{K^+}^2 - m_{K^0}^2}{m_{K^0}^2 - m_{K^+}^2 + m_{\pi^+}^2} = 0.56, \\ \frac{m_s}{m_d} &= \frac{m_{K^0}^2 + m_{K^+}^2 - m_{\pi^+}^2}{m_{K^0}^2 + m_{\pi^+}^2 - m_{K^+}^2} = 20.2, \end{aligned} \quad (10)$$

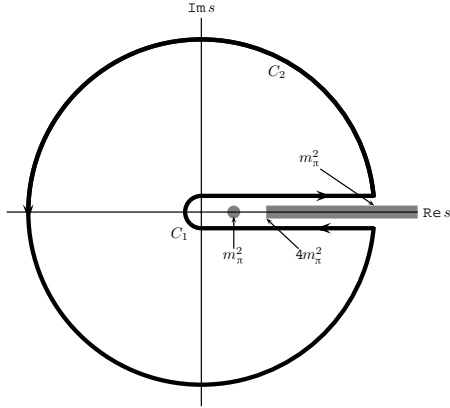
to lowest order in chiral perturbation theory, with an error which will be estimated below. Since the mass ratios extracted using chiral perturbation theory use the symmetry transformation property of  $M$  under the chiral symmetry  $G_\chi$ , it is important to use a renormalization scheme for QCD that does not change this transformation law. Any mass independent subtraction scheme such as  $\overline{\text{MS}}$  is suitable. The ratios of quark masses are scale independent in such a scheme, and Eq. (10) can be taken to be the ratio of  $\overline{\text{MS}}$  masses. Chiral perturbation theory cannot determine the overall scale of the quark masses, since it uses only the symmetry properties of  $M$ , and any multiple of  $M$  has the same  $G_\chi$  transformation law as  $M$ .

Chiral perturbation theory is a systematic expansion in powers of the light quark masses. The typical expansion parameter is  $m_K^2/\Lambda_\chi^2 \sim 0.25$  if one uses  $SU(3)$  chiral symmetry, and  $m_\pi^2/\Lambda_\chi^2 \sim 0.02$  if instead one uses  $SU(2)$  chiral symmetry. Electromagnetic effects at the few percent level also break

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$SU(2)$  and  $SU(3)$  symmetry. The mass formulae Eq. (9) were derived using  $SU(3)$  chiral symmetry, and are expected to have approximately a 25% uncertainty due to second order corrections. This estimate of the uncertainty is consistent with the lattice results found in Eq. (3)-Eq. (5) and more recent calculations.



**Figure 1:** The analytic structure of  $\Pi(s)$  in the complex  $s$ -plane. The contours  $C_1$  and  $C_2$  are the integration contours discussed in the text.

There is a subtlety which arises when one tries to determine quark mass ratios at second order in chiral perturbation theory. The second order quark mass term [10]

$$\left(M^\dagger\right)^{-1} \det M^\dagger \quad (11)$$

(which can be generated by instantons) transforms in the same way under  $G_\chi$  as  $M$ . Chiral perturbation theory cannot distinguish between  $M$  and  $\left(M^\dagger\right)^{-1} \det M^\dagger$ ; one can make the replacement  $M \rightarrow M(\lambda) = M + \lambda M \left(M^\dagger M\right)^{-1} \det M^\dagger$  in the chiral Lagrangian,

$$\begin{aligned} M(\lambda) &= \text{diag}(m_u(\lambda), m_d(\lambda), m_s(\lambda)) \\ &= \text{diag}(m_u + \lambda m_d m_s, m_d + \lambda m_u m_s, m_s + \lambda m_u m_d), \end{aligned} \quad (12)$$

and leave all observables unchanged.

The combination

$$\left(\frac{m_u}{m_d}\right)^2 + \frac{1}{Q^2} \left(\frac{m_s}{m_d}\right)^2 = 1 \quad (13)$$

where

$$Q^2 = \frac{m_s^2 - \hat{m}^2}{m_d^2 - m_u^2}, \quad \hat{m} = \frac{1}{2}(m_u + m_d),$$

is insensitive to the transformation in Eq. (12). Eq. (13) gives an ellipse in the  $m_u/m_d - m_s/m_d$  plane. The ellipse is well-determined by chiral perturbation theory, but the exact location on the ellipse, and the absolute normalization of the quark masses, has large uncertainties.  $Q$  is determined to be

in the range 21–25 from  $\eta \rightarrow 3\pi$  decay and the electromagnetic contribution to the  $K^+ - K^0$  and  $\pi^+ - \pi^0$  mass differences [11].

The absolute normalization of the quark masses cannot be determined using chiral perturbation theory. Other methods, such as lattice simulations discussed above or spectral function sum rules [12,13] for hadronic correlation functions, which we review next are necessary.

**Sum Rules:** Sum rule methods have been used extensively to determine quark masses and for illustration we briefly discuss here their application to hadronic  $\tau$  decays [14]. Other applications involve very similar techniques.

The experimentally measured quantity is  $R_\tau$ ,

$$\frac{dR_\tau}{ds} = \frac{d\Gamma/ds(\tau^- \rightarrow \text{hadrons} + \nu_\tau(\gamma))}{\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau(\gamma))} \quad (14)$$

the hadronic invariant mass spectrum in semihadronic  $\tau$  decay, normalized to the leptonic  $\tau$  decay rate. It is useful to define  $q$  as the total momentum of the hadronic final state, so  $s = q^2$  is the hadronic invariant mass. The total hadronic  $\tau$  decay rate  $R_\tau$  is then given by integrating  $dR_\tau/ds$  over the kinematically allowed range  $0 \leq s \leq M_\tau^2$ .

$R_\tau$  can be written as

$$\begin{aligned} R_\tau &= 12\pi \int_0^{M_\tau^2} \frac{ds}{M_\tau^2} \left(1 - \frac{s}{M_\tau^2}\right)^2 \\ &\times \left[ \left(1 + 2\frac{s}{M_\tau^2}\right) \text{Im} \Pi^T(s) + \text{Im} \Pi^L(s) \right] \end{aligned} \quad (15)$$

where  $s = q^2$ , and the hadronic spectral functions  $\Pi^{L,T}$  are defined from the time-ordered correlation function of two weak currents is the time-ordered correlator of the weak interaction current ( $j^\mu(x)$  and  $j^\nu(0)$ ) by

$$\Pi^{\mu\nu}(q) = i \int d^4x e^{iq \cdot x} \langle 0 | T \left( j^\mu(x) j^\nu(0)^\dagger \right) | 0 \rangle, \quad (16)$$

$$\Pi^{\mu\nu}(q) = (-g^{\mu\nu} + q^\mu q^\nu) \Pi^T(s) + q^\mu q^\nu \Pi^L(s), \quad (17)$$

and the decomposition Eq. (17) is the most general possible structure consistent with Lorentz invariance.

By the optical theorem, the imaginary part of  $\Pi^{\mu\nu}$  is proportional to the total cross-section for the current to produce all possible states. A detailed analysis including the phase space factors leads to Eq. (15). The spectral functions  $\Pi^{L,T}(s)$  are analytic in the complex  $s$  plane, with singularities along the real axis. There is an isolated pole at  $s = m_\pi^2$ , and single- and multi-particle singularities for  $s \geq 4m_\pi^2$ , the two-particle threshold. The discontinuity along the real axis is  $\Pi^{L,T}(s + i0^+) - \Pi^{L,T}(s - i0^+) = 2i \text{Im} \Pi^{L,T}(s)$ . As a result, Eq. (15) can be rewritten with the replacement  $\text{Im} \Pi^{L,T}(s) \rightarrow -i \Pi^{L,T}(s)/2$ , and the integration being over the contour  $C_1$ . Finally, the contour  $C_1$  can be deformed to  $C_2$  without crossing any singularities, and so leaving the integral unchanged. One can derive a series of sum rules analogous to Eq. (15) by

weighting the differential  $\tau$  hadronic decay rate by different powers of the hadronic invariant mass,

$$R_{\tau}^{kl} = \int_0^{M_{\tau}^2} ds \left(1 - \frac{s}{M_{\tau}^2}\right)^k \left(\frac{s}{M_{\tau}^2}\right)^l \frac{dR_{\tau}}{ds} \quad (18)$$

where  $dR_{\tau}/ds$  is the hadronic invariant mass distribution in  $\tau$  decay normalized to the leptonic decay rate. This leads to the final form of the sum rule(s),

$$R_{\tau}^{kl} = -6\pi i \int_{C_2} \frac{ds}{M_{\tau}^2} \left(1 - \frac{s}{M_{\tau}^2}\right)^{2+k} \left(\frac{s}{M_{\tau}^2}\right)^l \times \left[ \left(1 + 2\frac{s}{M_{\tau}^2}\right) \Pi^T(s) + \Pi^L(s) \right]. \quad (19)$$

The manipulations so far are completely rigorous and exact, relying only on the general analytic structure of quantum field theory. The left-hand side of the sum rule Eq. (19) is obtained from experiment. The right hand-side can be computed for  $s$  far away from any physical cuts using the operator product expansion (OPE) for the time-ordered product of currents in Eq. (16), and QCD perturbation theory. The OPE is an expansion for the time-ordered product Eq. (16) in a series of local operators, and is an expansion about the  $q \rightarrow \infty$  limit. It gives  $\Pi(s)$  as an expansion in powers of  $\alpha_s(s)$  and  $\Lambda_{\text{QCD}}^2/s$ , and is valid when  $s$  is far (in units of  $\Lambda_{\text{QCD}}^2$ ) from any singularities in the complex  $s$ -plane.

The OPE gives  $\Pi(s)$  as a series in  $\alpha_s$ , quark masses, and various non-perturbative vacuum matrix element. By computing  $\Pi(s)$  theoretically, and comparing with the experimental values of  $R_{\tau}^{kl}$ , one determines various parameters such as  $\alpha_s$  and the quark masses. The theoretical uncertainties in using Eq. (19) arise from neglected higher order corrections (both perturbative and non-perturbative), and because the OPE is no longer valid near the real axis, where  $\Pi$  has singularities. The contribution of neglected higher order corrections can be estimated as for any other perturbative computation. The error due to the failure of the OPE is more difficult to estimate. In Eq. (19), the OPE fails on the endpoints of  $C_2$  that touch the real axis at  $s = M_{\tau}^2$ . The weight factor  $(1 - s/M_{\tau}^2)$  in Eq. (19) vanishes at this point, so the importance of the endpoint can be reduced by choosing larger values of  $k$ .

### E. Heavy quarks

For heavy-quark physics one can exploit the fact that  $m_Q \gg \Lambda_{\text{QCD}}$  to construct effective theories ( $m_Q$  is the mass of the heavy quark  $Q$ ). The masses and decay rates of hadrons containing a single heavy quark, such as the  $B$  and  $D$  mesons can be determined using the heavy quark effective theory (HQET) [45]. The theoretical calculations involve radiative corrections computed in perturbation theory with an expansion in  $\alpha_s(m_Q)$  and non-perturbative corrections with an expansion in powers of  $\Lambda_{\text{QCD}}/m_Q$ . Due to the asymptotic nature of the QCD perturbation series, the two kinds of corrections are intimately related; an example of this are renormalon effects

in the perturbative expansion which are associated with non-perturbative corrections.

Systems containing two heavy quarks such as the  $\Upsilon$  or  $J/\Psi$  are treated using non-relativistic QCD (NRQCD) [46]. The typical momentum and energy transfers in these systems are  $\alpha_s m_Q$ , and  $\alpha_s^2 m_Q$ , respectively, so these bound states are sensitive to scales much smaller than  $m_Q$ . However, smeared observables, such as the cross-section for  $e^+e^- \rightarrow b\bar{b}$  averaged over some range of  $s$  that includes several bound state energy levels, are better behaved and only sensitive to scales near  $m_Q$ . For this reason, most determinations of the  $c, b$  quark masses using perturbative calculations compare smeared observables with experiment [47–49].

There are many continuum extractions of the  $c$  and  $b$  quark masses, some with quoted errors of 10 MeV or smaller. There are systematic effects of comparable size, which are typically not included in these error estimates. Reference [41], for example, shows that even though the error estimate of  $m_c$  using the rapid convergence of the  $\alpha_s$  perturbation series is only a few MeV, the central value of  $m_c$  can differ by a much larger amount depending on which algorithm (all of which are formally equally good) is used to determine  $m_c$  from the data. This leads to a systematic error from perturbation theory of around 20 MeV for the  $c$  quark and 25 MeV for the  $b$  quark. Electromagnetic effects, which also are important at this precision, are often not included. For this reason, we inflate the errors on the continuum extractions of  $m_c$  and  $m_b$ . The average values of  $m_c$  and  $m_b$  from continuum determinations are (see Sec. G for the 1S scheme)

$$\overline{m}_c(\overline{m}_c) = (1.28 \pm 0.025) \text{ GeV}$$

$$\overline{m}_b(\overline{m}_b) = (4.18 \pm 0.03) \text{ GeV}, \quad m_b^{1S} = (4.65 \pm 0.03) \text{ GeV}.$$

Lattice simulations of QCD lead to discretization errors which are powers of  $m_Q a$  (modulated by logarithms); the power depends on the formulation of lattice QCD being used and in most cases is quadratic. Clearly these errors can be reduced by performing simulations at smaller lattice spacings, but also by using *improved* discretizations of the theory. Recently, with more powerful computing resources, better algorithms and techniques, it has become possible to perform simulations in the charm quark region and beyond, also decreasing the extrapolation which has to be performed to reach the  $b$ -quark. A novel approach proposed in [64] has been to compare the lattice results for moments of correlation functions of  $c\bar{c}$  quark-bilinear operators to perturbative calculations of the same quantities at 4-loop order. In this way both the strong coupling constant and the charm quark mass can be determined with remarkably small errors; in particular  $\overline{m}_c(\overline{m}_c) = 1.273(6) \text{ GeV}$  [36]. This lattice determination also uses the perturbative expression for the current-current correlator, and so has the perturbation theory systematic error discussed above. Recent updates using this correlator method, both with a very similar result, can be found in [27,37]. It should be remembered that these results

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were obtained in QCD with exact isospin symmetry; isospin breaking effects, including electromagnetism may well be larger or of the order of the quoted uncertainty.

As the range of heavy-quark masses which can be used in numerical simulations increases, results obtained by extrapolating the results to  $b$ -physics are becoming ever more reliable (see e.g. [27]). Traditionally however, the main approach to controlling the discretization errors in lattice studies of heavy quark physics has been to perform simulations of the effective theories such as HQET and NRQCD. This remains an important technique, both in its own right and in providing additional information for extrapolations from lower masses to the bottom region. Using effective theories,  $m_b$  is obtained from what is essentially a computation of the difference of  $M_{H_b} - m_b$ , where  $M_{H_b}$  is the mass of a hadron  $H_b$  containing a  $b$ -quark. The relative error on  $m_b$  is therefore much smaller than that for  $M_{H_b} - m_b$ . The principal systematic errors are the matching of the effective theories to QCD and the presence of power divergences in  $a^{-1}$  in the  $1/m_b$  corrections which have to be subtracted numerically. The use of HQET or NRQCD is less precise for the charm quark, but in this case, as mentioned above, direct QCD simulations are now possible.

### F. Pole Mass

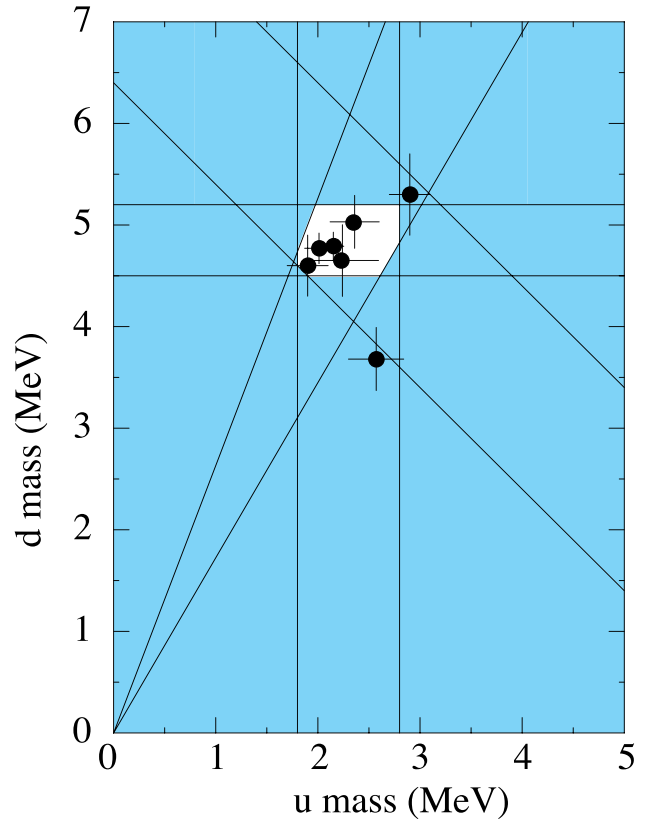
For an observable particle such as the electron, the position of the pole in the propagator is the definition of its mass. In QCD this definition of the quark mass is known as the pole mass. It is known that the on-shell quark propagator has no infrared divergences in perturbation theory [52,53], so this provides a perturbative definition of the quark mass. The pole mass cannot be used to arbitrarily high accuracy because of nonperturbative infrared effects in QCD. The full quark propagator has no pole because the quarks are confined, so that the pole mass cannot be defined outside of perturbation theory. The relation between the pole mass  $m_Q$  and the  $\overline{\text{MS}}$  mass  $\overline{m}_Q$  is known to three loops [54,55,56,57]

$$m_Q = \overline{m}_Q(\overline{m}_Q) \left\{ 1 + \frac{4\overline{\alpha}_s(\overline{m}_Q)}{3\pi} + \left[ -1.0414 \sum_k \left( 1 - \frac{4}{3} \frac{\overline{m}_{Q_k}}{\overline{m}_Q} \right) + 13.4434 \right] \left[ \frac{\overline{\alpha}_s(\overline{m}_Q)}{\pi} \right]^2 + [0.6527N_L^2 - 26.655N_L + 190.595] \left[ \frac{\overline{\alpha}_s(\overline{m}_Q)}{\pi} \right]^3 \right\}, \quad (20)$$

where  $\overline{\alpha}_s(\mu)$  is the strong interaction coupling constants in the  $\overline{\text{MS}}$  scheme, and the sum over  $k$  extends over the  $N_L$  flavors  $Q_k$  lighter than  $Q$ . The complete mass dependence of the  $\alpha_s^2$  term can be found in [54]; the mass dependence of the  $\alpha_s^3$  term is not known. For the  $b$ -quark, Eq. (20) reads

$$m_b = \overline{m}_b(\overline{m}_b) [1 + 0.10 + 0.05 + 0.03], \quad (21)$$

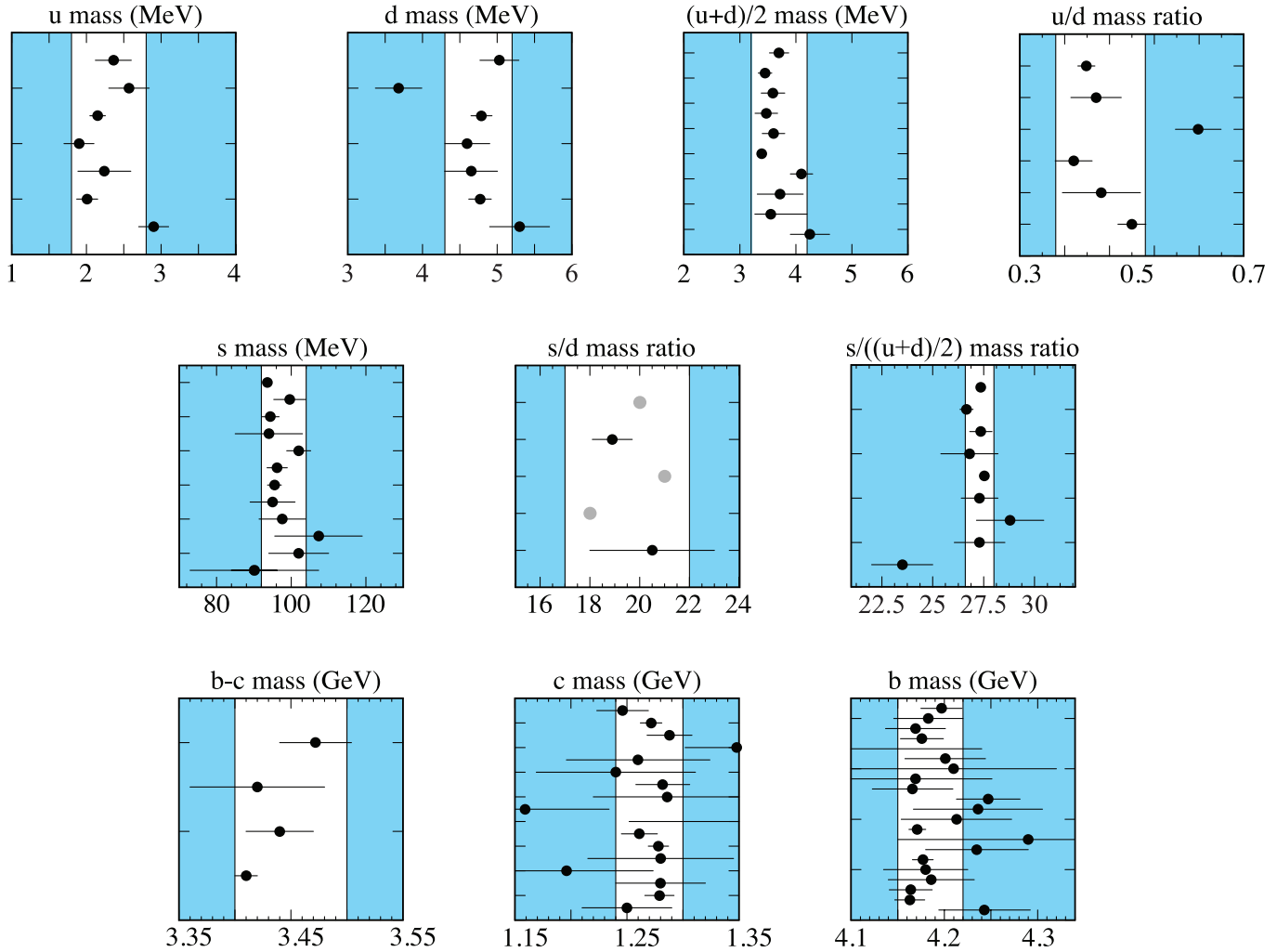
where the contributions from the different orders in  $\alpha_s$  are shown explicitly. The two and three loop corrections are comparable in size and have the same sign as the one loop term. This is a signal of the asymptotic nature of the perturbation series [there is a renormalon in the pole mass]. Such a badly behaved perturbation expansion can be avoided by directly extracting the  $\overline{\text{MS}}$  mass from data without extracting the pole mass as an intermediate step.



**Figure 2:** The allowed region (shown in white) for up quark and down quark masses. This region was determined in part from papers reporting values for  $m_u$  and  $m_d$  (data points shown) and in part from analysis of the allowed ranges of other mass parameters (see Fig. 3). The parameter  $(m_u + m_d)/2$  yields the two downward-sloping lines, while  $m_u/m_d$  yields the two rising lines originating at  $(0,0)$ .

### G. Numerical values and caveats

The quark masses in the particle data listings have been obtained by using a wide variety of methods. Each method involves its own set of approximations and uncertainties. In most cases, the errors are an estimate of the size of neglected higher-order corrections or other uncertainties. The expansion parameters for some of the approximations are not very small (for example, they are  $m_K^2/\Lambda_\chi^2 \sim 0.25$  for the chiral expansion



**Figure 3.** The values of each quark mass parameter taken from the Data Listings. The points are in chronological order with the more recent measurements at the top. Points from papers reporting no error bars are colored grey. The shaded regions indicate values excluded by our evaluations; some regions were determined in part through examination of Fig. 2.

and  $\Lambda_{\text{QCD}}/m_b \sim 0.1$  for the heavy-quark expansion), so an unexpectedly large coefficient in a neglected higher-order term could significantly alter the results. It is also important to note that the quark mass values can be significantly different in the different schemes.

The heavy quark masses obtained using HQET, QCD sum rules, or lattice gauge theory are consistent with each other if they are all converted into the same scheme and scale. We have specified all masses in the  $\overline{\text{MS}}$  scheme. For light quarks, the renormalization scale has been chosen to be  $\mu = 2 \text{ GeV}$ . The light quark masses at 1 GeV are significantly different from those at 2 GeV,  $\overline{m}(1 \text{ GeV})/\overline{m}(2 \text{ GeV}) \sim 1.33$ . It is conventional to choose the renormalization scale equal to the quark mass for

a heavy quark, so we have quoted  $\overline{m}_Q(\mu)$  at  $\mu = \overline{m}_Q$  for the  $c$  and  $b$  quarks. Recent analyses of inclusive  $B$  meson decays have shown that recently proposed mass definitions lead to a better behaved perturbation series than for the  $\overline{\text{MS}}$  mass, and hence to more accurate mass values. We have chosen to also give values for one of these, the  $b$  quark mass in the 1S-scheme [58,59]. Other schemes that have been proposed are the PS-scheme [60] and the kinetic scheme [61].

If necessary, we have converted values in the original papers to our chosen scheme using two-loop formulae. It is important to realize that our conversions introduce significant additional errors. In converting to the  $\overline{\text{MS}}$   $b$ -quark mass, for example, the three-loop conversions from the 1S and pole masses give values about 35 MeV and 135 MeV lower than the two-loop

# Quark Particle Listings

## Quarks, $u$

conversions. The uncertainty in  $\alpha_s(M_Z) = 0.1181(13)$  gives an uncertainty of  $\pm 10$  MeV and  $\pm 35$  MeV respectively in the same conversions. We have not added these additional errors when we do our conversions. The  $\alpha_s$  value in the conversion is correlated with the  $\alpha_s$  value used in determining the quark mass, so the conversion error is not a simple additional error on the quark mass.

### References

1. See the review of QCD in this volume..
2. A.V. Manohar and H. Georgi, Nucl. Phys. **B234**, 189 (1984).
3. K.G. Chetyrkin, Phys. Lett. **B404**, 161 (1997).
4. J.A.M. Vermaseren, S.A. Larin, and T. van Ritbergen, Phys. Lett. **B405**, 327 (1997).
5. K.G. Chetyrkin, B.A. Kniehl, and M. Steinhauser, Nucl. Phys. **B510**, 61 (1998).
6. S. Weinberg, Physica **96A**, 327 (1979).
7. J. Gasser and H. Leutwyler, Ann. Phys. **158**, 142 (1984).
8. For a review, see A. Pich, Rept. on Prog. in Phys. **58**, 563 (1995).
9. S. Weinberg, Trans. N.Y. Acad. Sci. **38**, 185 (1977).
10. D.B. Kaplan and A.V. Manohar, Phys. Rev. Lett. **56**, 2004 (1986).
11. H. Leutwyler, Phys. Lett. **B374**, 163 (1996).
12. S. Weinberg, Phys. Rev. Lett. **18**, 507 (1967).
13. M.A. Shifman, A.I. Vainshtein, and V.I. Zakharov, Nucl. Phys. **B147**, 385 (1979).
14. E. Braaten, S. Narison, and A. Pich, Nucl. Phys. **B373**, 581 (1992).
15. C. Bernard *et al.*, PoS **LAT2007** (2007) 090.
16. S. Aoki *et al.* [FLAG Collab.], Eur. Phys. J. **C74**, 2890 (2014).
17. A. Bazavov *et al.*, arXiv:0903.3598 [hep-lat].
18. C. Aubin *et al.* [HPQCD Collab.], Phys. Rev. **D70**, 031504 (2004).
19. C. Aubin *et al.* [MILC Collab.], Phys. Rev. **D70**, 114501 (2004).
20. B. Blossier *et al.* [ETM Collab.], Phys. Rev. **D82**, 114513 (2010).
21. A. Bazavov *et al.* [MILC Collab.], PoS **CD09** (2009) 007.
22. A. Bazavov *et al.*, PoS **LATTICE2010** (2010) 083.
23. S. Durr *et al.*, Phys. Lett. **B701**, 265 (2011).
24. S. Durr *et al.*, J. High Energy Phys. **1108**, 148 (2011).
25. R. Arthur *et al.* [RBC and UKQCD Collabs.], Phys. Rev. **D87**, 094514 (2013).
26. A. Bazavov *et al.* [Fermilab Lattice and MILC Collabs.], Phys. Rev. **D90**, 074509 (2014).
27. B. Chakraborty *et al.*, Phys. Rev. **D91**, 054508 (2015).
28. N. Carrasco *et al.* [European Twisted Mass Collab.], Nucl. Phys. **B887**, 19 (2014).
29. "Domain wall QCD with physical quark masses," T. Blum *et al.* [RBC and UKQCD Collabs.], arXiv:1411.7017 [hep-lat].
30. S. Borsanyi *et al.*, Science **347**, 1452 (2015).
31. Y. Aoki *et al.* [RBC and UKQCD Collabs.], Phys. Rev. **D83**, 074508 (2011).
32. S. Basak *et al.* [MILC Collab.], J. Phys. Conf. Ser. **640** (2015) 1, 012052.
33. T. Blum *et al.*, Phys. Rev. **D82**, 094508 (2010).
34. S. Aoki *et al.*, Phys. Rev. **D86**, 034507 (2012).
35. C.T.H. Davies *et al.*, Phys. Rev. Lett. **104**, 132003 (2010).
36. C. McNeile *et al.*, Phys. Rev. **D82**, 034512 (2010).
37. K. Nakayama, B. Fahy, and S. Hashimoto, arXiv:1511.09163 [hep-lat].
38. C. Aubin *et al.* [MILC Collab.], Nucl. Phys. (Proc. Supp.) **140**, 231 (2005).
39. C. Aubin *et al.* [MILC Collab.], Phys. Rev. **D70**, 114501 (2004).
40. G. Colangelo *et al.*, Eur. Phys. J. **C71**, 1695 (2011).
41. B. Dehnadi *et al.*, arXiv:1102.2264 [hep-ph].
42. T. Blum *et al.*, Phys. Rev. **D76**, 114508 (2007).
43. G. Colangelo *et al.*, Eur. Phys. J. **C71**, 1695 (2011).
44. A. Ali Khan *et al.* [CP-PACS Collab.], Phys. Rev. **D65**, 054505 (2002); [Erratum-ibid. **D67** (2003) 059901].
45. N. Isgur and M.B. Wise, Phys. Lett. **B232**, 113 (1989), *ibid.*, **B237**, 527 (1990).
46. G.T. Bodwin, E. Braaten, and G.P. Lepage, Phys. Rev. **D51**, 1125 (1995).
47. A.H. Hoang, Phys. Rev. **D61**, 034005 (2000).
48. K. Melnikov and A. Yelkhovsky, Phys. Rev. **D59**, 114009 (1999).
49. M. Beneke and A. Signer, Phys. Lett. **B471**, 233 (1999).
50. A.X. El-Khadra, A.S. Kronfeld, and P.B. Mackenzie, Phys. Rev. **D55**, 3933 (1997).
51. S. Aoki, Y. Kuramashi, and S.i. Tominaga, Prog. Theor. Phys. **109**, 383 (2003).
52. R. Tarrach, Nucl. Phys. **B183**, 384 (1981).
53. A. Kronfeld, Phys. Rev. **D58**, 051501 (1998).
54. N. Gray *et al.*, Z. Phys. **C48**, 673 (1990).
55. D.J. Broadhurst, N. Gray, and K. Schilcher, Z. Phys. **C52**, 111 (1991).
56. K.G. Chetyrkin and M. Steinhauser, Phys. Rev. Lett. **83**, 4001 (1999).
57. K. Melnikov and T. van Ritbergen, Phys. Lett. **B482**, 99 (2000).
58. A.H. Hoang, Z. Ligeti, A.V. Manohar, Phys. Rev. Lett. **82**, 277 (1999).
59. A.H. Hoang, Z. Ligeti, A.V. Manohar, Phys. Rev. **D59**, 074017 (1999).
60. M. Beneke, Phys. Lett. **B434**, 115 (1998).
61. P. Gambino and N. Uraltsev, Eur. Phys. J. **C34**, 181 (2004).
62. G. Martinelli *et al.*, Nucl. Phys. **B445**, 81 (1995).
63. K. Jansen *et al.*, Phys. Lett. **B372**, 275 (1996).
64. I. Allison *et al.* [HPQCD Collab.], Phys. Rev. **D78**, 054513 (2008).

**$u$**

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$$

$$\text{Mass } m = 2.2^{+0.6}_{-0.4} \text{ MeV}$$

$$\text{Charge} = \frac{2}{3} e$$

$$I_z = +\frac{1}{2}$$

$$m_u/m_d = 0.38\text{--}0.58$$



See key on page 601

# Quark Particle Listings

## $d, s$ , Light Quarks ( $u, d, s$ )

<b>d</b>	$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$	
Mass $m = 4.7^{+0.5}_{-0.4}$ MeV	Charge $= -\frac{1}{3} e$	$I_z = -\frac{1}{2}$
$m_s/m_d = 17\text{--}22$		
$\overline{m} = (m_u + m_d)/2 = 3.5^{+0.7}_{-0.3}$ MeV		

<b>S</b>	$I(J^P) = 0(\frac{1}{2}^+)$	
Mass $m = 96^{+8}_{-4}$ MeV	Charge $= -\frac{1}{3} e$	Strangeness $= -1$
$(m_s - (m_u + m_d)/2)/(m_d - m_u) = 27.3 \pm 0.7$		

## Light Quarks ( $u, d, s$ )

OMITTED FROM SUMMARY TABLE

### $u$ -QUARK MASS

The  $u$ -,  $d$ -, and  $s$ -quark masses are estimates of so-called “current-quark masses,” in a mass-independent subtraction scheme such as  $\overline{\text{MS}}$ . The ratios  $m_u/m_d$  and  $m_s/m_d$  are extracted from pion and kaon masses using chiral symmetry. The estimates of  $d$  and  $u$  masses are not without controversy and remain under active investigation. Within the literature there are even suggestions that the  $u$  quark could be essentially massless. The  $s$ -quark mass is estimated from SU(3) splittings in hadron masses.

We have normalized the  $\overline{\text{MS}}$  masses at a renormalization scale of  $\mu = 2$  GeV. Results quoted in the literature at  $\mu = 1$  GeV have been rescaled by dividing by 1.35. The values of “Our Evaluation” were determined in part via Figures 1 and 2.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2.2 <math>\pm 0.6</math> -0.4</b> OUR EVALUATION	See the ideogram below.		
2.36 $\pm 0.24$	1 CARRASCO	14 LATT	$\overline{\text{MS}}$ scheme
2.57 $\pm 0.26 \pm 0.07$	2 AOKI	12 LATT	$\overline{\text{MS}}$ scheme
2.15 $\pm 0.03 \pm 0.10$	3 DURR	11 LATT	$\overline{\text{MS}}$ scheme
1.9 $\pm 0.2$	4 BAZAVOV	10 LATT	$\overline{\text{MS}}$ scheme
2.24 $\pm 0.10 \pm 0.34$	5 BLUM	10 LATT	$\overline{\text{MS}}$ scheme
2.01 $\pm 0.14$	6 MCNEILE	10 LATT	$\overline{\text{MS}}$ scheme
2.9 $\pm 0.2$	7 DOMINGUEZ	09 THEO	$\overline{\text{MS}}$ scheme
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.01 $\pm 0.14$	6 DAVIES	10 LATT	$\overline{\text{MS}}$ scheme
2.9 $\pm 0.8$	8 DEANDREA	08 THEO	$\overline{\text{MS}}$ scheme
3.02 $\pm 0.33$	9 BLUM	07 LATT	$\overline{\text{MS}}$ scheme
2.7 $\pm 0.4$	10 JAMIN	06 THEO	$\overline{\text{MS}}$ scheme
1.9 $\pm 0.2$	11 MASON	06 LATT	$\overline{\text{MS}}$ scheme
2.8 $\pm 0.2$	12 NARISON	06 THEO	$\overline{\text{MS}}$ scheme
1.7 $\pm 0.3$	13 AUBIN	04A LATT	$\overline{\text{MS}}$ scheme

1 CARRASCO 14 is a lattice QCD computation of light quark masses using  $2 + 1 + 1$  dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.

2 AOKI 12 is a lattice computation using  $1 + 1 + 1$  dynamical quark flavors.

3 DURR 11 determine quark mass from a lattice computation of the meson spectrum using  $N_f = 2 + 1$  dynamical flavors. The lattice simulations were done at the physical quark mass, so that extrapolation in the quark mass was not needed. The individual  $m_u, m_d$  values are obtained using the lattice determination of the average mass  $m_{ud}$  and of the ratio  $m_s/m_{ud}$  and the value of  $Q = (m_s^2 - m_{ud}^2) / (m_d^2 - m_u^2)$  as determined from  $\eta \rightarrow 3\pi$  decays.

4 BAZAVOV 10 is a lattice computation using  $2+1$  dynamical quark flavors.

5 BLUM 10 determines light quark masses using a QCD plus QED lattice computation of the electromagnetic mass splittings of the low-lying hadrons. The lattice simulations use  $2+1$  dynamical quark flavors.

6 DAVIES 10 and MCNEILE 10 determine  $\overline{m}_c(\mu)/\overline{m}_s(\mu) = 11.85 \pm 0.16$  using a lattice computation with  $N_f = 2 + 1$  dynamical fermions of the pseudoscalar meson masses. Mass  $m_p$  is obtained from this using the value of  $m_c$  from ALLISON 08 or MCNEILE 10 and the BAZAVOV 10 values for the light quark mass ratios,  $m_s/\overline{m}$  and  $m_u/m_d$ .

7 DOMINGUEZ 09 use QCD finite energy sum rules for the two-point function of the divergence of the axial vector current computed to order  $\alpha_s^4$ .

8 DEANDREA 08 determine  $m_u - m_d$  from  $\eta \rightarrow 3\pi^0$ , and combine with the PDG 06 lattice average value of  $m_u + m_d = 7.6 \pm 1.6$  to determine  $m_u$  and  $m_d$ .

9 BLUM 07 determine quark masses from the pseudoscalar meson masses using a QED plus QCD lattice computation with two dynamical quark flavors.

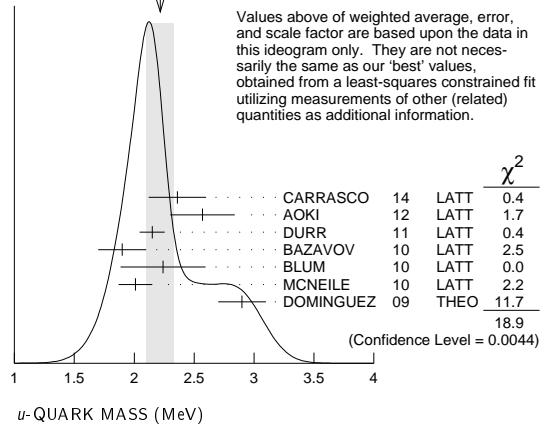
10 JAMIN 06 determine  $m_u(2 \text{ GeV})$  by combining the value of  $m_s$  obtained from the spectral function for the scalar  $K\pi$  form factor with other determinations of the quark mass ratios.

11 MASON 06 extract light quark masses from a lattice simulation using staggered fermions with an improved action, and three dynamical light quark flavors with degenerate  $u$  and  $d$  quarks. Perturbative corrections were included at NNLO order. The quark masses  $m_u$  and  $m_d$  were determined from their  $(m_u + m_d)/2$  measurement and AUBIN 04A  $m_u/m_d$  value.

12 NARISON 06 uses sum rules for  $e^+e^- \rightarrow$  hadrons to order  $\alpha_s^3$  to determine  $m_s$  combined with other determinations of the quark mass ratios.

13 AUBIN 04A employ a partially quenched lattice calculation of the pseudoscalar meson masses.

WEIGHTED AVERAGE  
2.22  $\pm 0.12$  (Error scaled by 1.8)



### $d$ -QUARK MASS

See the comment for the  $u$  quark above.

We have normalized the  $\overline{\text{MS}}$  masses at a renormalization scale of  $\mu = 2$  GeV. Results quoted in the literature at  $\mu = 1$  GeV have been rescaled by dividing by 1.35. The values of “Our Evaluation” were determined in part via Figures 1 and 2.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4.7 <math>\pm 0.5</math> -0.4</b> OUR EVALUATION	See the ideogram below.		
5.03 $\pm 0.26$	1 CARRASCO	14 LATT	$\overline{\text{MS}}$ scheme
3.68 $\pm 0.29 \pm 0.10$	2 AOKI	12 LATT	$\overline{\text{MS}}$ scheme
4.79 $\pm 0.07 \pm 0.12$	3 DURR	11 LATT	$\overline{\text{MS}}$ scheme
4.6 $\pm 0.3$	4 BAZAVOV	10 LATT	$\overline{\text{MS}}$ scheme
4.65 $\pm 0.15 \pm 0.32$	5 BLUM	10 LATT	$\overline{\text{MS}}$ scheme
4.77 $\pm 0.15$	6 MCNEILE	10 LATT	$\overline{\text{MS}}$ scheme
5.3 $\pm 0.4$	7 DOMINGUEZ	09 THEO	$\overline{\text{MS}}$ scheme
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.79 $\pm 0.16$	6 DAVIES	10 LATT	$\overline{\text{MS}}$ scheme
4.7 $\pm 0.8$	8 DEANDREA	08 THEO	$\overline{\text{MS}}$ scheme
5.49 $\pm 0.39$	9 BLUM	07 LATT	$\overline{\text{MS}}$ scheme
4.8 $\pm 0.5$	10 JAMIN	06 THEO	$\overline{\text{MS}}$ scheme
4.4 $\pm 0.3$	11 MASON	06 LATT	$\overline{\text{MS}}$ scheme
5.1 $\pm 0.4$	12 NARISON	06 THEO	$\overline{\text{MS}}$ scheme
3.9 $\pm 0.5$	13 AUBIN	04A LATT	$\overline{\text{MS}}$ scheme

1 CARRASCO 14 is a lattice QCD computation of light quark masses using  $2 + 1 + 1$  dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.

2 AOKI 12 is a lattice computation using  $1 + 1 + 1$  dynamical quark flavors.

3 DURR 11 determine quark mass from a lattice computation of the meson spectrum using  $N_f = 2 + 1$  dynamical flavors. The lattice simulations were done at the physical quark mass, so that extrapolation in the quark mass was not needed. The individual  $m_u, m_d$  values are obtained using the lattice determination of the average mass  $m_{ud}$  and of the ratio  $m_s/m_{ud}$  and the value of  $Q = (m_s^2 - m_{ud}^2) / (m_d^2 - m_u^2)$  as determined from  $\eta \rightarrow 3\pi$  decays.

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5 BLUM 10 determines light quark masses using a QCD plus QED lattice computation of the electromagnetic mass splittings of the low-lying hadrons. The lattice simulations use  $2+1$  dynamical quark flavors.

6 DAVIES 10 and MCNEILE 10 determine  $\overline{m}_c(\mu)/\overline{m}_s(\mu) = 11.85 \pm 0.16$  using a lattice computation with  $N_f = 2 + 1$  dynamical fermions of the pseudoscalar meson masses. Mass  $m_d$  is obtained from this using the value of  $m_c$  from ALLISON 08 or MCNEILE 10 and the BAZAVOV 10 values for the light quark mass ratios,  $m_s/\overline{m}$  and  $m_u/m_d$ .

7 DOMINGUEZ 09 use QCD finite energy sum rules for the two-point function of the divergence of the axial vector current computed to order  $\alpha_s^4$ .

8 DEANDREA 08 determine  $m_u - m_d$  from  $\eta \rightarrow 3\pi^0$ , and combine with the PDG 06 lattice average value of  $m_u + m_d = 7.6 \pm 1.6$  to determine  $m_u$  and  $m_d$ .

9 BLUM 07 determine quark masses from the pseudoscalar meson masses using a QED plus QCD lattice computation with two dynamical quark flavors.

10 JAMIN 06 determine  $m_d(2 \text{ GeV})$  by combining the value of  $m_s$  obtained from the spectral function for the scalar  $K\pi$  form factor with other determinations of the quark mass ratios.

11 MASON 06 extract light quark masses from a lattice simulation using staggered fermions with an improved action, and three dynamical light quark flavors with degenerate  $u$  and  $d$  quarks. Perturbative corrections were included at NNLO order. The quark masses  $m_u$  and  $m_d$  were determined from their  $(m_u + m_d)/2$  measurement and AUBIN 04A  $m_u/m_d$  value.

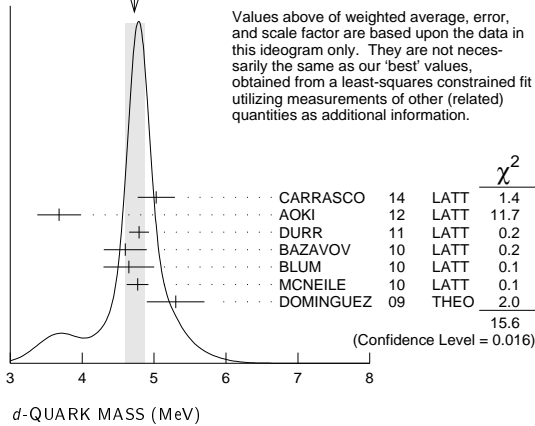
12 NARISON 06 uses sum rules for  $e^+e^- \rightarrow$  hadrons to order  $\alpha_s^3$  to determine  $m_s$  combined with other determinations of the quark mass ratios.

13 AUBIN 04A perform three flavor dynamical lattice calculation of pseudoscalar meson masses, with continuum estimate of electromagnetic effects in the kaon masses, and one-loop perturbative renormalization constant.

# Quark Particle Listings

## Light Quarks ( $u, d, s$ )

WEIGHTED AVERAGE  
4.73±0.13 (Error scaled by 1.6)



$$\bar{m} = (m_u + m_d)/2$$

See the comments for the  $u$  quark above.

We have normalized the  $\bar{M}$  masses at a renormalization scale of  $\mu = 2$  GeV. Results quoted in the literature at  $\mu = 1$  GeV have been rescaled by dividing by 1.35. The values of "Our Evaluation" were determined in part via Figures 1 and 2.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>3.5 ±0.3</b>	<b>OUR EVALUATION</b> See the ideogram below.		
3.70 ±0.17	1 CARRASCO	14 LATT	$\bar{M}$ scheme
3.45 ±0.12	2 ARTHUR	13 LATT	$\bar{M}$ scheme
3.59 ±0.21	3 AOKI	11A LATT	$\bar{M}$ scheme
3.469±0.047±0.048	4 DURR	11 LATT	$\bar{M}$ scheme
3.6 ±0.2	5 BLOSSIER	10 LATT	$\bar{M}$ scheme
3.39 ±0.06	6 MCNEILE	10 LATT	$\bar{M}$ scheme
4.1 ±0.2	7 DOMINGUEZ	09 THEO	$\bar{M}$ scheme
3.72 ±0.41	8 ALLTON	08 LATT	$\bar{M}$ scheme
3.55 +0.65 -0.28	9 ISHIKAWA	08 LATT	$\bar{M}$ scheme
4.25 ±0.35	10 BLUM	07 LATT	$\bar{M}$ scheme
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.40 ±0.07	6 DAVIES	10 LATT	$\bar{M}$ scheme
3.85 ±0.12 ±0.4	11 BLOSSIER	08 LATT	$\bar{M}$ scheme
≥ 4.85 ±0.20	12 DOMINGUEZ...08B	THEO	$\bar{M}$ scheme
4.026±0.048	13 NAKAMURA	08 LATT	$\bar{M}$ scheme
4.08 ±0.25 ±0.42	14 GOCKELER	06 LATT	$\bar{M}$ scheme
4.7 ±0.2 ±0.3	15 GOCKELER	06A LATT	$\bar{M}$ scheme
3.2 ±0.3	16 MASON	06 LATT	$\bar{M}$ scheme
3.95 ±0.3	17 NARISON	06 THEO	$\bar{M}$ scheme
2.8 ±0.3	18 AUBIN	04 LATT	$\bar{M}$ scheme
4.29 ±0.14 ±0.65	19 AOKI	03 LATT	$\bar{M}$ scheme
3.223±0.3	20 AOKI	03B LATT	$\bar{M}$ scheme
4.4 ±0.1 ±0.4	21 BECIREVIC	03 LATT	$\bar{M}$ scheme
4.1 ±0.3 ±1.0	22 CHIU	03 LATT	$\bar{M}$ scheme

1 CARRASCO 14 is a lattice QCD computation of light quark masses using  $2 + 1 + 1$  dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.

2 ARTHUR 13 is a lattice computation using  $2+1$  dynamical domain wall fermions. Masses at  $\mu = 3$  GeV have been converted to  $\mu = 2$  GeV using conversion factors given in their paper.

3 AOKI 11A determine quark masses from a lattice computation of the hadron spectrum using  $N_f = 2 + 1$  dynamical flavors of domain wall fermions.

4 DURR 11 determine quark mass from a lattice computation of the meson spectrum using  $N_f = 2 + 1$  dynamical flavors. The lattice simulations were done at the physical quark mass, so that extrapolation in the quark mass was not needed.

5 BLOSSIER 10 determines quark masses from a computation of the hadron spectrum using  $N_f=2$  dynamical twisted-mass Wilson fermions.

6 DAVIES 10 and MCNEILE 10 determine  $\bar{m}_C(\mu)/\bar{m}_S(\mu) = 11.85 \pm 0.16$  using a lattice computation with  $N_f = 2 + 1$  dynamical fermions of the pseudoscalar meson masses. Mass  $\bar{m}$  is obtained from this using the value of  $m_C$  from ALLISON 08 or MCNEILE 10 and the BAZAVOV 10 values for the light quark mass ratio,  $m_s/\bar{m}$ .

7 DOMINGUEZ 09 use QCD finite energy sum rules for the two-point function of the divergence of the axial vector current computed to order  $\alpha_s^4$ .

8 ALLTON 08 use a lattice computation of the  $\pi$ ,  $K$ , and  $\Omega$  masses with  $2+1$  dynamical flavors of domain wall quarks, and non-perturbative renormalization.

9 ISHIKAWA 08 use a lattice computation of the light meson spectrum with  $2+1$  dynamical flavors of  $\mathcal{O}(a)$  improved Wilson quarks, and one-loop perturbative renormalization.

10 BLUM 07 determine quark masses from the pseudoscalar meson masses using a QED plus QCD lattice computation with two dynamical quark flavors.

11 BLOSSIER 08 use a lattice computation of pseudoscalar meson masses and decay constants with 2 dynamical flavors and non-perturbative renormalization.

12 DOMINGUEZ-CLARIMON 08B obtain an inequality from sum rules for the scalar two-point correlator.

13 NAKAMURA 08 do a lattice computation using quenched domain wall fermions and non-perturbative renormalization.

14 GOCKELER 06 use an unquenched lattice computation of the axial Ward Identity with  $N_f = 2$  dynamical light quark flavors, and non-perturbative renormalization, to obtain  $\bar{m}(2 \text{ GeV}) = 4.08 \pm 0.25 \pm 0.19 \pm 0.23 \text{ MeV}$ , where the first error is statistical, the second and third are systematic due to the fit range and force scale uncertainties, respectively. We have combined the systematic errors linearly.

15 GOCKELER 06A use an unquenched lattice computation of the pseudoscalar meson masses with  $N_f = 2$  dynamical light quark flavors, and non-perturbative renormalization.

16 MASON 06 extract light quark masses from a lattice simulation using staggered fermions with an improved action, and three dynamical light quark flavors with degenerate  $u$  and  $d$  quarks. Perturbative corrections were included at NNLO order.

17 NARISON 06 uses sum rules for  $e^+e^- \rightarrow \text{hadrons}$  to order  $\alpha_s^3$  to determine  $m_s$  combined with other determinations of the quark mass ratios.

18 AUBIN 04 perform three flavor dynamical lattice calculation of pseudoscalar meson masses, with one-loop perturbative renormalization constant.

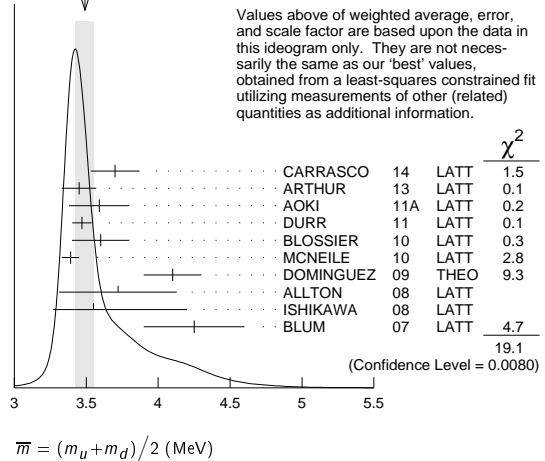
19 AOKI 03 uses quenched lattice simulation of the meson and baryon masses with degenerate light quarks. The extrapolations are done using quenched chiral perturbation theory.

20 The errors given in AOKI 03B were  $+0.046$   
 $-0.069$ . We changed them to  $\pm 0.3$  for calculating the overall best values. AOKI 03B uses lattice simulation of the meson and baryon masses with two dynamical light quarks. Simulations are performed using the  $\mathcal{O}(a)$  improved Wilson action.

21 BECIREVIC 03 perform quenched lattice computation using the vector and axial Ward identities. Uses  $\mathcal{O}(a)$  improved Wilson action and nonperturbative renormalization.

22 CHIU 03 determines quark masses from the pion and kaon masses using a lattice simulation with a chiral fermion action in quenched approximation.

WEIGHTED AVERAGE  
3.49±0.06 (Error scaled by 1.7)



$$\bar{m} = (m_u + m_d)/2 \text{ (MeV)}$$

### $m_u/m_d$ MASS RATIO

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.38-0.58</b>	<b>OUR EVALUATION</b> See the ideogram below.		
0.4482±0.0173 -0.0206	1 BASAK	15 LATT	
0.470 ±0.056	2 CARRASCO	14 LATT	
0.698 ±0.051	3 AOKI	12 LATT	
0.42 ±0.01 ±0.04	4 BAZAVOV	10 LATT	
0.4818±0.0096±0.0860	5 BLUM	10 LATT	
0.550 ±0.031	6 BLUM	07 LATT	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.43 ±0.08	7 AUBIN	04A LATT	
0.410 ±0.036	8 NELSON	03 LATT	
0.553 ±0.043	9 LEUTWYLER	96 THEO	Compilation

1 BASAK 15 is a lattice computation using  $2+1$  dynamical quark flavors.

2 CARRASCO 14 is a lattice QCD computation of light quark masses using  $2 + 1 + 1$  dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.

3 AOKI 12 is a lattice computation using  $1 + 1 + 1$  dynamical quark flavors.

4 BAZAVOV 10 is a lattice computation using  $2+1$  dynamical quark flavors.

5 BLUM 10 is a lattice computation using  $2+1$  dynamical quark flavors.

6 BLUM 07 determine quark masses from the pseudoscalar meson masses using a QED plus QCD lattice computation with two dynamical quark flavors.

7 AUBIN 04A perform three flavor dynamical lattice calculation of pseudoscalar meson masses, with continuum estimate of electromagnetic effects in the kaon masses.

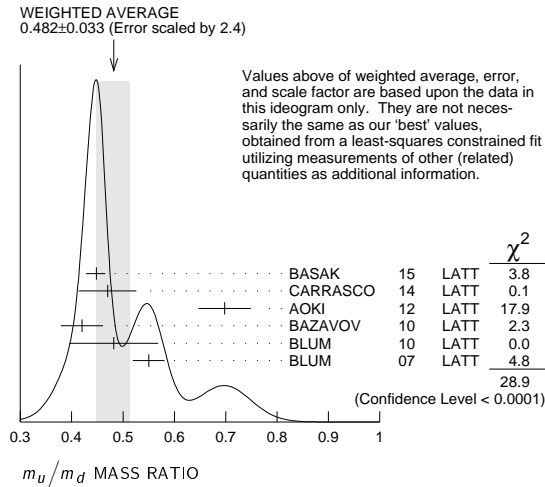
8 NELSON 03 computes coefficients in the order  $p^4$  chiral Lagrangian using a lattice calculation with three dynamical flavors. The ratio  $m_u/m_d$  is obtained by combining this with the chiral perturbation theory computation of the meson masses to order  $p^4$ .

See key on page 601

## Quark Particle Listings

### Light Quarks ( $u, d, s$ )

<sup>9</sup> LEUTWYLER 96 uses a combined fit to  $\eta \rightarrow 3\pi$  and  $\psi' \rightarrow J/\psi(\pi, \eta)$  decay rates, and the electromagnetic mass differences of the  $\pi$  and  $K$ .



#### s-QUARK MASS

See the comment for the  $u$  quark above.

We have normalized the  $\overline{MS}$  masses at a renormalization scale of  $\mu = 2$  GeV. Results quoted in the literature at  $\mu = 1$  GeV have been rescaled by dividing by 1.35.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>96 ± 8</b>	<b>OUR EVALUATION</b>		
93.6 ± 0.8	1 CHAKRABOR..15	LATT	$\overline{MS}$ scheme
99.6 ± 4.3	2 CARRASCO 14	LATT	$\overline{MS}$ scheme
94.4 ± 2.3	3 ARTHUR 13	LATT	$\overline{MS}$ scheme
94 ± 9	4 BODENSTEIN 13	THEO	$\overline{MS}$ scheme
102 ± 3 ± 1	5 FRITZSCH 12	LATT	$\overline{MS}$ scheme
96.2 ± 2.7	6 AOKI 11A	LATT	$\overline{MS}$ scheme
95.5 ± 1.1 ± 1.5	7 DURR 11	LATT	$\overline{MS}$ scheme
95 ± 6	8 BLOSSIER 10	LATT	$\overline{MS}$ scheme
97.6 ± 2.9 ± 5.5	9 BLUM 10	LATT	$\overline{MS}$ scheme
107.3 ± 11.7	10 ALLTON 08	LATT	$\overline{MS}$ scheme
102 ± 8	11 DOMINGUEZ 08A	THEO	$\overline{MS}$ scheme
90.1 ± 17.2	12 ISHIKAWA 08	LATT	$\overline{MS}$ scheme
90.1 ± 6.1			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
92.4 ± 1.5	13 DAVIES 10	LATT	$\overline{MS}$ scheme
92.2 ± 1.3	13 MCNEILE 10	LATT	$\overline{MS}$ scheme
105 ± 3 ± 9	14 BLOSSIER 08	LATT	$\overline{MS}$ scheme
105.6 ± 1.2	15 NAKAMURA 08	LATT	$\overline{MS}$ scheme
119.5 ± 9.3	16 BLUM 07	LATT	$\overline{MS}$ scheme
105 ± 6 ± 7	17 CHETYSKIN 06	THEO	$\overline{MS}$ scheme
111 ± 6 ± 10	18 GOKKELER 06	LATT	$\overline{MS}$ scheme
119 ± 5 ± 8	19 GOKKELER 06A	LATT	$\overline{MS}$ scheme
92 ± 9	20 JAMIN 06	THEO	$\overline{MS}$ scheme
87 ± 6	21 MASON 06	LATT	$\overline{MS}$ scheme
104 ± 15	22 NARISON 06	THEO	$\overline{MS}$ scheme
≥ 71 ± 4, ≤ 151 ± 14	23 NARISON 06	THEO	$\overline{MS}$ scheme
96 ± 5 ± 16	24 BAIKOV 05	THEO	$\overline{MS}$ scheme
81 ± 22	25 GAMIZ 05	THEO	$\overline{MS}$ scheme
125 ± 28	26 GORBUNOV 05	THEO	$\overline{MS}$ scheme
93 ± 32	27 NARISON 05	THEO	$\overline{MS}$ scheme
76 ± 8	28 AUBIN 04	LATT	$\overline{MS}$ scheme
116 ± 6 ± 0.65	29 AOKI 03	LATT	$\overline{MS}$ scheme
84.5 ± 12	30 AOKI 03B	LATT	$\overline{MS}$ scheme
106 ± 2 ± 8	31 BECIREVIC 03	LATT	$\overline{MS}$ scheme
92 ± 9 ± 16	32 CHIU 03	LATT	$\overline{MS}$ scheme
117 ± 17	33 GAMIZ 03	THEO	$\overline{MS}$ scheme
103 ± 17	34 GAMIZ 03	THEO	$\overline{MS}$ scheme

<sup>1</sup> CHAKRABORTY 15 is a lattice QCD computation that determines  $m_c$  and  $m_c/m_s$  using pseudoscalar meson masses tuned on gluon field configurations with 2+1+1 dynamical flavors of HISQ quarks with  $u/d$  masses down to the physical value.

<sup>2</sup> CARRASCO 14 is a lattice QCD computation of light quark masses using 2 + 1 + 1 dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.

<sup>3</sup> ARTHUR 13 is a lattice computation using 2+1 dynamical domain wall fermions. Masses at  $\mu = 3$  GeV have been converted to  $\mu = 2$  GeV using conversion factors given in their paper.

<sup>4</sup> BODENSTEIN 13 determines  $m_s$  from QCD finite energy sum rules, and the perturbative computation of the pseudoscalar correlator to five-loop order.

<sup>5</sup> FRITZSCH 12 determine  $m_s$  using a lattice computation with  $N_f = 2$  dynamical flavors.

<sup>6</sup> AOKI 11A determine quark masses from a lattice computation of the hadron spectrum using  $N_f = 2 + 1$  dynamical flavors of domain wall fermions.

<sup>7</sup> DURR 11 determine quark mass from a lattice computation of the meson spectrum using  $N_f = 2 + 1$  dynamical flavors. The lattice simulations were done at the physical quark mass, so that extrapolation in the quark mass was not needed.

<sup>8</sup> BLOSSIER 10 determines quark masses from a computation of the hadron spectrum using  $N_f = 2$  dynamical twisted-mass Wilson fermions.

<sup>9</sup> BLUM 10 determines light quark masses using a QCD plus QED lattice computation of the electromagnetic mass splittings of the low-lying hadrons. The lattice simulations use 2+1 dynamical quark flavors.

<sup>10</sup> ALLTON 08 use a lattice computation of the  $\pi$ ,  $K$ , and  $\Omega$  masses with 2+1 dynamical flavors of domain wall quarks, and non-perturbative renormalization.

<sup>11</sup> DOMINGUEZ 08A make determination from QCD finite energy sum rules for the pseudoscalar two-point function computed to order  $\alpha_s^4$ .

<sup>12</sup> ISHIKAWA 08 use a lattice computation of the light meson spectrum with 2+1 dynamical flavors of  $\mathcal{O}(a)$  improved Wilson quarks, and one-loop perturbative renormalization.

<sup>13</sup> DAVIES 10 and MCNEILE 10 determine  $\overline{m}_c(\mu)/\overline{m}_s(\mu) = 11.85 \pm 0.16$  using a lattice computation with  $N_f = 2 + 1$  dynamical fermions of the pseudoscalar meson masses. Mass  $m_c$  is obtained from this using the value of  $m_c$  from ALLISON 08 or MCNEILE 10.

<sup>14</sup> BLOSSIER 08 use a lattice computation of pseudoscalar meson masses and decay constants with 2 dynamical flavors and non-perturbative renormalization.

<sup>15</sup> NAKAMURA 08 do a lattice computation using quenched domain wall fermions and non-perturbative renormalization.

<sup>16</sup> BLUM 07 determine quark masses from the pseudoscalar meson masses using a QED plus QCD lattice computation with two dynamical quark flavors.

<sup>17</sup> CHETYSKIN 06 use QCD sum rules in the pseudoscalar channel to order  $\alpha_s^4$ .

<sup>18</sup> GOKKELER 06 use an unquenched lattice computation of the axial Ward Identity with  $N_f = 2$  dynamical light quark flavors, and non-perturbative renormalization, to obtain  $\overline{m}_s(2 \text{ GeV}) = 111 \pm 6 \pm 4 \pm 6 \text{ MeV}$ , where the first error is statistical, the second and third are systematic due to the fit range and force scale uncertainties, respectively. We have combined the systematic errors linearly.

<sup>19</sup> GOKKELER 06A use an unquenched lattice computation of the pseudoscalar meson masses with  $N_f = 2$  dynamical light quark flavors, and non-perturbative renormalization.

<sup>20</sup> JAMIN 06 determine  $\overline{m}_s(2 \text{ GeV})$  from the spectral function for the scalar  $K\pi$  form factor.

<sup>21</sup> MASON 06 extract light quark masses from a lattice simulation using staggered fermions with an improved action, and three dynamical light quark flavors with degenerate  $u$  and  $d$  quarks. Perturbative corrections were included at NNLO order.

<sup>22</sup> NARISON 06 uses sum rules for  $e^+e^- \rightarrow \text{hadrons}$  to order  $\alpha_s^3$ .

<sup>23</sup> NARISON 06 obtains the quoted range from positivity of the spectral functions.

<sup>24</sup> BAIKOV 05 determines  $\overline{m}_s(M_\tau) = 100^{+5}_{-3} + 17$  from sum rules using the strange spectral function in  $\tau$  decay. The computations were done to order  $\alpha_s^3$ , with an estimate of the  $\alpha_s^4$  terms. We have converted the result to  $\mu = 2$  GeV.

<sup>25</sup> GAMIZ 05 determines  $\overline{m}_s(2 \text{ GeV})$  from sum rules using the strange spectral function in  $\tau$  decay. The computations were done to order  $\alpha_s^2$ , with an estimate of the  $\alpha_s^3$  terms.

<sup>26</sup> GORBUNOV 05 use hadronic tau decays to N<sup>3</sup>LO, including power corrections.

<sup>27</sup> NARISON 05 determines  $\overline{m}_s(2 \text{ GeV})$  from sum rules using the strange spectral function in  $\tau$  decay. The computations were done to order  $\alpha_s^3$ .

<sup>28</sup> AUBIN 04 perform three flavor dynamical lattice calculation of pseudoscalar meson masses, with one-loop perturbative renormalization constant.

<sup>29</sup> AOKI 03 uses quenched lattice simulation of the meson and baryon masses with degenerate light quarks. The extrapolations are done using quenched chiral perturbation theory. Determines  $m_s = 113.8 \pm 2.3^{+5.8}_{-2.9}$  using  $K$  mass as input and  $m_s = 142.3 \pm 5.8^{+22}_{-0}$  using  $\phi$  mass as input. We have performed a weighted average of these values.

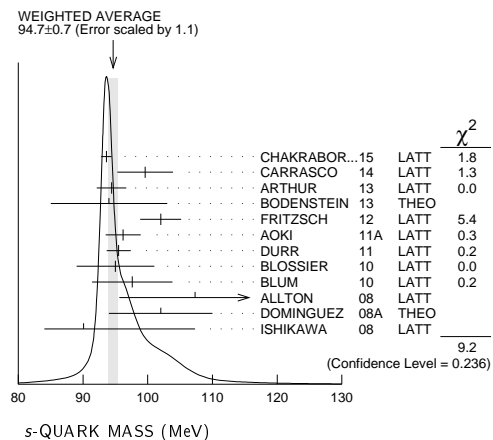
<sup>30</sup> AOKI 03B uses lattice simulation of the meson and baryon masses with two dynamical light quarks. Simulations are performed using the  $\mathcal{O}(a)$  improved Wilson action.

<sup>31</sup> BECIREVIC 03 perform quenched lattice computation using the vector and axial Ward identities. Uses  $\mathcal{O}(a)$  improved Wilson action and nonperturbative renormalization. They also quote  $\overline{m}/m_s = 24.3 \pm 0.2 \pm 0.6$ .

<sup>32</sup> CHIU 03 determines quark masses from the pion and kaon masses using a lattice simulation with a chiral fermion action in quenched approximation.

<sup>33</sup> GAMIZ 03 determines  $m_s$  from SU(3) breaking in the  $\tau$  hadronic width. The value of  $V_{us}$  is chosen to satisfy CKM unitarity.

<sup>34</sup> GAMIZ 03 determines  $m_s$  from SU(3) breaking in the  $\tau$  hadronic width. The value of  $V_{us}$  is taken from the PDG.



# Quark Particle Listings

## Light Quarks ( $u, d, s$ ), $c$

### OTHER LIGHT QUARK MASS RATIOS

#### $m_s/m_d$ MASS RATIO

VALUE	DOCUMENT ID	TECN	COMMENT
<b>17-22 OUR EVALUATION</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
20.0	1 GAO 97	THEO	
18.9±0.8	2 LEUTWYLER 96	THEO	Compilation
21	3 DONOGHUE 92	THEO	
18	4 GERARD 90	THEO	
18 to 23	5 LEUTWYLER 90b	THEO	

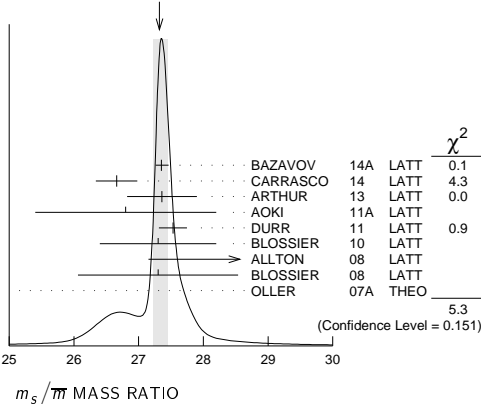
- 1 GAO 97 uses electromagnetic mass splittings of light mesons.
- 2 LEUTWYLER 96 uses a combined fit to  $\eta \rightarrow 3\pi$  and  $\psi' \rightarrow J/\psi(\pi, \eta)$  decay rates, and the electromagnetic mass differences of the  $\pi$  and  $K$ .
- 3 DONOGHUE 92 result is from a combined analysis of meson masses,  $\eta \rightarrow 3\pi$  using second-order chiral perturbation theory including nonanalytic terms, and  $(\psi(2S) \rightarrow J/\psi(1S)\pi)/(\psi(2S) \rightarrow J/\psi(1S)\eta)$ .
- 4 GERARD 90 uses large  $N$  and  $\eta$ - $\eta'$  mixing.
- 5 LEUTWYLER 90b determines quark mass ratios using second-order chiral perturbation theory for the meson and baryon masses, including nonanalytic corrections. Also uses Weinberg sum rules to determine  $L_7$ .

#### $m_s/\overline{m}$ MASS RATIO

VALUE	DOCUMENT ID	TECN
<b>27.3 ±0.7 OUR EVALUATION</b>		
See the ideogram below.		
27.35±0.05 <sup>+0.10</sup> <sub>-0.07</sub>	1 BAZAVOV 14A	LATT
26.66±0.32	2 CARRASCO 14	LATT
27.36±0.54	3 ARTHUR 13	LATT
26.8 ±1.4	4 AOKI 11A	LATT
27.53±0.20±0.08	5 DURR 11	LATT
27.3 ±0.9	6 BLOSSIER 10	LATT
28.8 ±1.65	7 ALLTON 08	LATT
27.3 ±0.3 ±1.2	8 BLOSSIER 08	LATT
23.5 ±1.5	9 OLLER 07A	THEO
27.4 ±0.4	10 AUBIN 04	LATT

- • • We do not use the following data for averages, fits, limits, etc. • • •
- 1 BAZAVOV 14A is a lattice computation using 4 dynamical flavors of HISQ fermions.
- 2 CARRASCO 14 is a lattice QCD computation of light quark masses using 2 + 1 + 1 dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.
- 3 ARTHUR 13 is a lattice computation using 2+1 dynamical domain wall fermions.
- 4 AOKI 11A determine quark masses from a lattice computation of the hadron spectrum using  $N_f = 2 + 1$  dynamical flavors of domain wall fermions.
- 5 DURR 11 determine quark mass from a lattice computation of the meson spectrum using  $N_f = 2 + 1$  dynamical flavors. The lattice simulations were done at the physical quark mass, so that extrapolation in the quark mass was not needed.
- 6 BLOSSIER 10 determines quark masses from a computation of the hadron spectrum using  $N_f=2$  dynamical twisted-mass Wilson fermions.
- 7 ALLTON 08 use a lattice computation of the  $\pi$ ,  $K$ , and  $\Omega$  masses with 2+1 dynamical flavors of domain wall quarks, and non-perturbative renormalization.
- 8 BLOSSIER 08 use a lattice computation of pseudoscalar meson masses and decay constants with 2 dynamical flavors and non-perturbative renormalization.
- 9 OLLER 07A use unitarized chiral perturbation theory to order  $p^4$ .
- 10 Three flavor dynamical lattice calculation of pseudoscalar meson masses.

WEIGHTED AVERAGE  
27.32±0.12-0.10 (Error scaled by 1.3)



#### Q MASS RATIO

$Q \equiv \sqrt{(m^2_s - \overline{m}^2)/(m^2_d - m^2_u)}$ ; $\overline{m} \equiv (m_u + m_d)/2$			
VALUE	DOCUMENT ID	TECN	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
22.8±0.4	1 MARTEMYA... 05	THEO	
22.7±0.8	2 ANISOVICH 96	THEO	

- 1 MARTEMYANOV 05 determine  $Q$  from  $\eta \rightarrow 3\pi$  decay.
- 2 ANISOVICH 96 find  $Q$  from  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decay using dispersion relations and chiral perturbation theory.

### LIGHT QUARKS ( $u, d, s$ ) REFERENCES

BASAK 15	JPCS 640 012052	S. Basak <i>et al.</i>	(MILC Collab.)
CHAKRABOR... 15	PR D91 054508	B. Chakraborty <i>et al.</i>	(HPQCD Collab.)
BAZAVOV 14A	PR D90 074509	A. Bazavov <i>et al.</i>	(Fermi-LAT and MILC Collabs.)
CARRASCO 14	NP B887 19	N. Carrasco <i>et al.</i>	(European Twisted Mass Collab.)
ARTHUR 13	PR D87 094514	R. Arthur <i>et al.</i>	(RBC and UKQCD Collabs.)
BODENSTEIN 13	JHEP 1307 138	S. Bodenstein, C.A. Dominguez, K. Schilcher (MANZ+)	
AOKI 12	PR D86 034507	S. Aoki <i>et al.</i>	(PACS-CS Collab.)
FRITZSCH 12	NP B865 397	P. Fritzsch <i>et al.</i>	(ALPHA Collab.)
AOKI 11A	PR D83 074508	Y. Aoki <i>et al.</i>	(RBC-UKQCD Collab.)
DURR 11	PL B701 265	S. Durr <i>et al.</i>	(BMW Collab.)
BAZAVOV 10	RMP 82 1349	A. Bazavov <i>et al.</i>	(MILC Collab.)
BLOSSIER 10	PR D82 114513	B. Blossier <i>et al.</i>	(ETM Collab.)
BLUM 10	PR D82 094508	T. Blum <i>et al.</i>	
DAVIES 10	PRL 104 132003	C.T.H. Davies <i>et al.</i>	(HPQCD Collab.)
MCNEILE 10	PR D82 034512	C. McNeile <i>et al.</i>	(HPQCD Collab.)
DOMINGUEZ 09	PR D79 014009	C.A. Dominguez <i>et al.</i>	
ALLISON 08	PR D78 054513	I. Allison <i>et al.</i>	(HPQCD Collab.)
ALLTON 08	PR D78 114509	C. Allton <i>et al.</i>	(RBC and UKQCD Collabs.)
BLOSSIER 08	JHEP 0804 020	B. Blossier <i>et al.</i>	(ETM Collab.)
DEANDREA 08	PR D78 034032	A. Deandrea, A. Nehme, P. Talavera	
DOMINGUEZ 08A	JHEP 0805 020	C.A. Dominguez <i>et al.</i>	
DOMINGUEZ... 08B	PL B660 49	A. Dominguez-Clarimon, E. de Rafael, J. Taron	
ISHIKAWA 08	PR D78 011502	T. Ishikawa <i>et al.</i>	(CP-PACS and JLQCD Collabs.)
NAKAMURA 08	PR D78 034502	Y. Nakamura <i>et al.</i>	(CP-PACS Collab.)
BLUM 07	PR D76 114508	T. Blum <i>et al.</i>	(RBC Collab.)
OLLER 07A	EPJ A34 371	J.A. Oller, L. Roca	
CHETYRKIN 06	EPJ C46 721	K.G. Chetyrkin, A. Khodjamirian	
GOCKELER 06	PR D73 054508	M. Gockeler <i>et al.</i>	(QCDSF, UKQCD Collabs.)
GOCKELER 06A	PL B639 307	M. Gockeler <i>et al.</i>	(QCDSF, UKQCD Collabs.)
JAMIN 06	PR D74 074009	M. Jamin, J.A. Oller, A. Pich	
MASON 06	PR D73 114501	Q. Mason <i>et al.</i>	(HPQCD Collab.)
NARISON 06	PR D74 034013	S. Narison	
PDG 06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
BAIKOV 05	PRL 95 012003	P.A. Baikov, K.G. Chetyrkin, J.H. Kuhn	
GAMIZ 05	PRL 94 011803	E. Gamiz <i>et al.</i>	
GORBUNOV 05	PR D71 013002	D.S. Gorbunov, A.A. Pivovarov	
MARTEMYA... 05	PR D71 017501	B.V. Martemyanov, V.S. Sopov	
NARISON 05	PL B626 101	S. Narison	
AUBIN 04	PR D70 031504	C. Aubin <i>et al.</i>	(HPQCD, MILC, UKQCD Collabs.)
AUBIN 04A	PR D70 114501	C. Aubin <i>et al.</i>	(MILC Collab.)
AOKI 03	PR D67 034503	S. Aoki <i>et al.</i>	(CP-PACS Collab.)
AOKI 03B	PR D68 054502	S. Aoki <i>et al.</i>	(CP-PACS Collab.)
BECEVIC 03	PL B558 69	D. Becirevic, V. Lubicz, C. Tarantino	
CHIU 03	NP B673 217	T.-W. Chiu, T.-H. Hsieh	
GAMIZ 03	JHEP 0301 060	E. Gamiz <i>et al.</i>	
NELSON 03	PRL 90 021601	D. Nelson, G.T. Fleming, G.W. Kilcup	
GAO 97	PR D56 4115	D.-N. Gao, B.A. Li, M.-L. Yan	
ANISOVICH 96	PL B375 335	A.V. Anisovich, H. Leutwyler	
LEUTWYLER 96	PL B378 313	H. Leutwyler	
DONOGHUE 92	PRL 69 3444	J.F. Donoghue, B.R. Holstein, D. Wyler	(MASA+)
GERARD 90	MPL A5 391	J.M. Gerard	(MPIM)
LEUTWYLER 90B	NP B337 108	H. Leutwyler	(BERN)



$I(J^P) = 0(\frac{1}{2}^+)$

Charge =  $\frac{2}{3} e$  Charm = +1

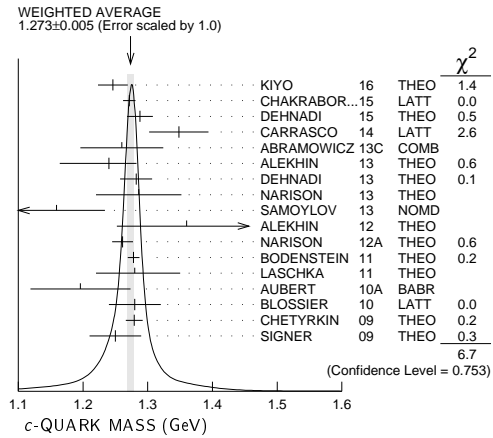
### c-QUARK MASS

The  $c$ -quark mass corresponds to the “running” mass  $m_c(\mu = m_c)$  in the  $\overline{MS}$  scheme. We have converted masses in other schemes to the  $\overline{MS}$  scheme using two-loop QCD perturbation theory with  $\alpha_s(\mu=m_c) = 0.38 \pm 0.03$ . The value  $1.27 \pm 0.03$  GeV for the  $\overline{MS}$  mass corresponds to  $1.67 \pm 0.07$  GeV for the pole mass (see the “Note on Quark Masses”).

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b>1.27 ±0.03 OUR EVALUATION</b>			
See the ideogram below.			
1.246 ±0.023	1 KIYO 16	THEO	$\overline{MS}$ scheme
1.2715 ±0.0095	2 CHAKRABOR... 15	LATT	$\overline{MS}$ scheme
1.288 ±0.020	3 DEHNADI 15	THEO	$\overline{MS}$ scheme
1.348 ±0.046	4 CARRASCO 14	LATT	$\overline{MS}$ scheme
1.26 ±0.05 ±0.04	5 ABRAMOWICZ13c	COMB	$\overline{MS}$ scheme
1.24 ±0.03 ±0.03 <sub>-0.07</sub>	6 ALEKHIN 13	THEO	$\overline{MS}$ scheme
1.282 ±0.011 ±0.022	7 DEHNADI 13	THEO	$\overline{MS}$ scheme
1.286 ±0.066	8 NARISON 13	THEO	$\overline{MS}$ scheme
1.159 ±0.075	9 SAMOYLOV 13	NOMD	$\overline{MS}$ scheme
1.36 ±0.04 ±0.10	10 ALEKHIN 12	THEO	$\overline{MS}$ scheme
1.261 ±0.016	11 NARISON 12A	THEO	$\overline{MS}$ scheme
1.278 ±0.009	12 BODENSTEIN 11	THEO	$\overline{MS}$ scheme
1.28 ±0.07 <sub>-0.06</sub>	13 LASCHKA 11	THEO	$\overline{MS}$ scheme
1.28 ±0.04	15 BLOSSIER 10	LATT	$\overline{MS}$ scheme
1.279 ±0.013	16 CHETYRKIN 09	THEO	$\overline{MS}$ scheme
1.25 ±0.04	17 SIGNER 09	THEO	$\overline{MS}$ scheme
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.01 ±0.09 ±0.03	18 ALEKHIN 11	THEO	$\overline{MS}$ scheme
1.299 ±0.026	19 BODENSTEIN 10	THEO	$\overline{MS}$ scheme
1.273 ±0.006	20 MCNEILE 10	LATT	$\overline{MS}$ scheme
1.261 ±0.018	21 NARISON 10	THEO	$\overline{MS}$ scheme
1.268 ±0.009	22 ALLISON 08	LATT	$\overline{MS}$ scheme
1.286 ±0.013	23 KUHN 07	THEO	$\overline{MS}$ scheme
1.295 ±0.015	24 BOUGHEZAL 06	THEO	$\overline{MS}$ scheme

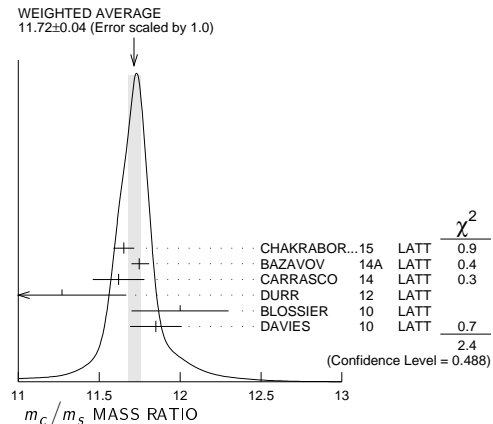
1.24 ± 0.09	25	BUCHMUEL...	06	THEO	$\overline{MS}$ scheme
1.224 ± 0.017 ± 0.054	26	HOANG	06	THEO	$\overline{MS}$ scheme
1.33 ± 0.10	27	AUBERT	04x	THEO	$\overline{MS}$ scheme
1.29 ± 0.07	28	HOANG	04	THEO	$\overline{MS}$ scheme
1.319 ± 0.028	29	DEDIVITIIS	03	LATT	$\overline{MS}$ scheme
1.19 ± 0.11	30	EIDEMULLER	03	THEO	$\overline{MS}$ scheme
1.289 ± 0.043	31	ERLER	03	THEO	$\overline{MS}$ scheme
1.26 ± 0.02	32	ZYABLYUK	03	THEO	$\overline{MS}$ scheme

- <sup>1</sup> KIYO 16 determine  $\overline{m}_c(\overline{m}_c)$  from the  $J/\psi(1S)$  mass at order  $\alpha_s^3$  (N3LO).
- <sup>2</sup> CHAKRABORTY 15 is a lattice QCD computation using 2+1+1 dynamical flavors. Moments of pseudoscalar current-current correlators are matched to  $\alpha_s^3$ -accurate QCD perturbation theory with the  $\eta_c$  meson mass tuned to experiment.
- <sup>3</sup> DEHNADI 15 determine  $\overline{m}_c(\overline{m}_c)$  using sum rules for  $e^+e^- \rightarrow$  hadrons at order  $\alpha_s^3$  (N3LO), and fitting to both experimental data and lattice results.
- <sup>4</sup> CARRASCO 14 is a lattice QCD computation of light quark masses using 2 + 1 + 1 dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.
- <sup>5</sup> ABRAMOWICZ 13c determines  $m_c$  from charm production in deep inelastic  $ep$  scattering, using the QCD prediction at NLO order. The uncertainties from model and parameterization assumptions, and the value of  $\alpha_s$ , of  $\pm 0.03$ ,  $\pm 0.02$ , and  $\pm 0.02$  respectively, have been combined in quadrature.
- <sup>6</sup> ALEKHIN 13 determines  $m_c$  from charm production in deep inelastic scattering at HERA using approximate NNLO QCD.
- <sup>7</sup> DEHNADI 13 determines  $m_c$  using QCD sum rules for the charmonium spectrum and charm continuum to order  $\alpha_s^3$  (N3LO). The statistical and systematic experimental errors of  $\pm 0.006$  and  $\pm 0.009$  have been combined in quadrature. The theoretical uncertainties  $\pm 0.019$  from truncation of the perturbation series,  $\pm 0.010$  from  $\alpha_s$ , and  $\pm 0.002$  from the gluon condensate have been combined in quadrature.
- <sup>8</sup> NARISON 13 determines  $m_c$  using QCD spectral sum rules to order  $\alpha_s^2$  (NNLO) and including condensates up to dimension 6.
- <sup>9</sup> SAMOYLOV 13 determines  $m_c$  from a study of charm dimuon production in neutrino-ion scattering using the NLO QCD result for the charm quark production cross section.
- <sup>10</sup> ALEKHIN 12 determines  $m_c$  from heavy quark production in deep inelastic scattering at HERA using approximate NNLO QCD.
- <sup>11</sup> NARISON 12A determines  $m_c$  using sum rules for the vector current correlator to order  $\alpha_s^3$ , including the effect of gluon condensates up to dimension eight.
- <sup>12</sup> BODENSTEIN 11 determine  $\overline{m}_c(3 \text{ GeV}) = 0.987 \pm 0.009 \text{ GeV}$  and  $\overline{m}_c(\overline{m}_c) = 1.278 \pm 0.009 \text{ GeV}$  using QCD sum rules for the charm quark vector current correlator.
- <sup>13</sup> LASCHKA 11 determine the  $c$  mass from the charmonium spectrum. The theoretical computation uses the heavy  $Q\overline{Q}$  potential to order  $1/m_Q$  obtained by matching the short-distance perturbative result onto lattice QCD result at larger scales.
- <sup>14</sup> AUBERT 10A determine the  $b$ - and  $c$ -quark masses from a fit to the inclusive decay spectra in semileptonic  $B$  decays in the kinetic scheme (and convert it to the  $\overline{MS}$  scheme).
- <sup>15</sup> BLOSSIER 10 determines quark masses from a computation of the hadron spectrum using  $N_f=2$  dynamical twisted-mass Wilson fermions.
- <sup>16</sup> CHETYRKIN 09 determine  $m_c$  and  $m_b$  from the  $e^+e^- \rightarrow Q\overline{Q}$  cross-section and sum rules, using an order  $\alpha_s^3$  computation of the heavy quark vacuum polarization. They also determine  $m_c(3 \text{ GeV}) = 0.986 \pm 0.013 \text{ GeV}$ .
- <sup>17</sup> SIGNER 09 determines the  $c$ -quark mass using non-relativistic sum rules to analyze the  $e^+e^- \rightarrow c\overline{c}$  cross-section near threshold. Also determine the PS mass  $m_{PS}(\mu_F = 0.7 \text{ GeV}) = 1.50 \pm 0.04 \text{ GeV}$ .
- <sup>18</sup> ALEKHIN 11 determines  $m_c$  from heavy quark production in deep inelastic scattering using fixed target and HERA data, and approximate NNLO QCD.
- <sup>19</sup> BODENSTEIN 10 determines  $\overline{m}_c(3 \text{ GeV}) = 1.008 \pm 0.026 \text{ GeV}$  using finite energy sum rules for the vector current correlator. The authors have converted this to  $\overline{m}_c(\overline{m}_c)$  using  $\alpha_s(M_Z) = 0.1189 \pm 0.0020$ .
- <sup>20</sup> MCNEILE 10 determines  $m_c$  by comparing the order  $\alpha_s^3$  perturbative results for the pseudo-scalar current to lattice simulations with  $N_f = 2+1$  sea-quarks by the HPQCD collaboration.
- <sup>21</sup> NARISON 10 determines  $m_c$  from ratios of moments of vector current correlators computed to order  $\alpha_s^3$  and including the dimension-six gluon condensate.
- <sup>22</sup> ALLISON 08 determine  $m_c$  by comparing four-loop perturbative results for the pseudo-scalar current correlator to lattice simulations by the HPQCD collaboration. The result has been updated in MCNEILE 10.
- <sup>23</sup> KUHN 07 determine  $\overline{m}_c(\mu = 3 \text{ GeV}) = 0.986 \pm 0.013 \text{ GeV}$  and  $\overline{m}_c(\overline{m}_c)$  from a four-loop sum-rule computation of the cross-section for  $e^+e^- \rightarrow$  hadrons in the charm threshold region.
- <sup>24</sup> BOUGHEZAL 06 result comes from the first moment of the hadronic production cross-section to order  $\alpha_s^3$ .
- <sup>25</sup> BUCHMUELLER 06 determine  $m_b$  and  $m_c$  by a global fit to inclusive  $B$  decay spectra.
- <sup>26</sup> HOANG 06 determines  $\overline{m}_c(\overline{m}_c)$  from a global fit to inclusive  $B$  decay data. The  $B$  decay distributions were computed to order  $\alpha_s^2\beta_0$ , and the conversion between different  $m_c$  mass schemes to order  $\alpha_s^3$ .
- <sup>27</sup> AUBERT 04x obtain  $m_c$  from a fit to the hadron mass and lepton energy distributions in semileptonic  $B$  decay. The paper quotes values in the kinetic scheme. The  $\overline{MS}$  value has been provided by the BABAR collaboration.
- <sup>28</sup> HOANG 04 determines  $\overline{m}_c(\overline{m}_c)$  from moments at order  $\alpha_s^2$  of the charm production cross-section in  $e^+e^-$  annihilation.
- <sup>29</sup> DEDIVITIIS 03 use a quenched lattice computation of heavy-heavy and heavy-light meson masses.
- <sup>30</sup> EIDEMULLER 03 determines  $m_b$  and  $m_c$  using QCD sum rules.
- <sup>31</sup> ERLER 03 determines  $m_b$  and  $m_c$  using QCD sum rules. Includes recent BES data.
- <sup>32</sup> ZYABLYUK 03 determines  $m_c$  by using QCD sum rules in the pseudoscalar channel and comparing with the  $\eta_c$  mass.

 **$m_c/m_s$  MASS RATIO**

VALUE	DOCUMENT ID	TECN
<b>11.72 ± 0.25 OUR EVALUATION</b>		See the ideogram below.
11.652 ± 0.065	<sup>1</sup> CHAKRABOR..15	LATT
11.747 ± 0.019 ± 0.059 -0.043	<sup>2</sup> BAZAVOV 14A	LATT
11.62 ± 0.16	<sup>3</sup> CARRASCO 14	LATT
11.27 ± 0.30 ± 0.26	<sup>4</sup> DURR 12	LATT
12.0 ± 0.3	<sup>5</sup> BLOSSIER 10	LATT
11.85 ± 0.16	<sup>6</sup> DAVIES 10	LATT

- <sup>1</sup> CHAKRABORTY 15 is a lattice QCD computation on gluon field configurations with 2+1+1 dynamical flavors of HISQ quarks with  $u/d$  masses down to the physical value.  $m_c$  and  $m_s$  are tuned from pseudoscalar meson masses.
- <sup>2</sup> BAZAVOV 14A is a lattice computation using 4 dynamical flavors of HISQ fermions.
- <sup>3</sup> CARRASCO 14 is a lattice QCD computation of light quark masses using 2 + 1 + 1 dynamical quarks, with  $m_u = m_d \neq m_s \neq m_c$ . The  $u$  and  $d$  quark masses are obtained separately by using the  $K$  meson mass splittings and lattice results for the electromagnetic contributions.
- <sup>4</sup> DURR 12 determine  $m_c/m_s$  using a lattice computation with  $N_f = 2$  dynamical fermions. The result is combined with other determinations of  $m_c$  to obtain  $m_s(2 \text{ GeV}) = 97.0 \pm 2.6 \pm 2.5 \text{ MeV}$ .
- <sup>5</sup> BLOSSIER 10 determine  $m_c/m_s$  from a computation of the hadron spectrum using  $N_f = 2$  dynamical twisted-mass Wilson fermions.
- <sup>6</sup> DAVIES 10 determine  $m_c/m_s$  from meson masses calculated on gluon fields including  $u$ ,  $d$ , and  $s$  sea quarks with lattice spacing down to 0.045 fm. The Highly Improved Staggered quark formalism is used for the valence quarks.

 **$m_b/m_c$  MASS RATIO**

VALUE	DOCUMENT ID	TECN
<b>4.528 ± 0.054</b>	<sup>1</sup> CHAKRABOR..15	LATT

- <sup>1</sup> CHAKRABORTY 15 is a lattice computation using 4 dynamical quark flavors.

 **$m_b - m_c$  QUARK MASS DIFFERENCE**

VALUE (GeV)	DOCUMENT ID	TECN
<b>3.45 ± 0.05 OUR EVALUATION</b>		
3.472 ± 0.032	<sup>1</sup> AUBERT 10A	BABR
3.42 ± 0.06	<sup>2</sup> ABDALLAH 06B	DLPH
3.44 ± 0.03	<sup>3</sup> AUBERT 04x	BABR
3.41 ± 0.01	<sup>3</sup> BAUER 04	THEO

- <sup>1</sup> AUBERT 10A determine the  $b$ - and  $c$ -quark masses from a fit to the inclusive decay spectra in semileptonic  $B$  decays in the kinetic scheme.
- <sup>2</sup> ABDALLAH 06B determine  $m_b - m_c$  from moments of the hadron invariant mass and lepton energy spectra in semileptonic inclusive  $B$  decays.
- <sup>3</sup> Determine  $m_b - m_c$  from a global fit to inclusive  $B$  decay spectra.

# Quark Particle Listings

c, b

## c-QUARK REFERENCES

KIYO	16	PL B752 122	Y. Kiyo, G. Mishima, Y. Sumino	
CHAKRABOR...	15	PR D91 054508	B. Chakraborty <i>et al.</i>	(HPQCD Collab.)
DEHNADI	15	JHEP 1508 155	B. Dehnadi, A.H. Hoang, V. Mateu	
BAZAVOV	14A	PR D90 074509	A. Bazavov <i>et al.</i>	(Fermi-LAT and MILC Collabs.)
CARRASCO	14	NP B887 19	N. Carrasco <i>et al.</i>	(European Twisted Mass Collab.)
ABRAMOWICZ	13C	EPJ C73 2311	H. Abramowicz <i>et al.</i>	(H1 and Zeus Collabs.)
ALEKHIN	13	PL B720 172	S. Alekhin <i>et al.</i>	(SERP, DESYZ, WUPP+)
DEHNADI	13	JHEP 1309 103	B. Dehnadi <i>et al.</i>	(SHRZ, VIEN, MPIM+)
NARISON	13	PL B718 1321	S. Narison	(MONP)
SAMOYLOV	13	NP B876 339	O. Samoylov <i>et al.</i>	(NOMAD Collab.)
ALEKHIN	12	PL B718 550	S. Alekhin <i>et al.</i>	(SERP, WUPP, DESY+)
DURR	12	PRL 108 122003	S. Durr, G. Koutsou	(WUPP, JULI, CYPR)
NARISON	12A	PL B706 412	S. Narison	(MONP)
ALEKHIN	11	PL B699 345	S. Alekhin, S. Moch	(DESY, SERP)
BODENSTEIN	11	PR D83 074014	S. Bodenstein <i>et al.</i>	
LASCHKA	11	PR D83 094002	A. Laschka, N. Kaiser, W. Weise	
AUBERT	10A	PR D81 032003	B. Aubert <i>et al.</i>	(BABAR Collab.)
BLOSSIER	10	PR D82 114513	B. Blossier <i>et al.</i>	(ETM Collab.)
BODENSTEIN	10	PR D82 114013	S. Bodenstein <i>et al.</i>	
DAVIES	10	PRL 104 132003	C.T.H. Davies <i>et al.</i>	(HPQCD Collab.)
MCNEILE	10	PR D82 034512	C. McNeile <i>et al.</i>	(HPQCD Collab.)
NARISON	10	PL B693 559	S. Narison	(MONP)
Also		PL B705 544 (err.)	S. Narison	(MONP)
CHETYRKIN	09	PR D80 074010	K.G. Chetyrkin <i>et al.</i>	(KARL, BNL)
SIGNER	09	PL B672 333	A. Signer	(DURH)
ALLISON	08	PR D78 054513	I. Allison <i>et al.</i>	(HPQCD Collab.)
KUHN	07	NP B778 192	J.H. Kuhn, M. Steinhauser, C. Sturm	
ABDALLAH	06B	EPJ C45 35	R. Abdallah <i>et al.</i>	(DELPHI Collab.)
BOUGHEZAL	06	PR D74 074006	R. Boughezal, M. Czakon, T. Schutzmeier	
BUCHMUEL...	06	PR D73 073008	O.L. Buchmueller, H.U. Flacher	(RHBL)
HOANG	05	PL B633 526	A.H. Hoang, A.V. Manohar	
AUBERT	04X	PRL 93 011803	B. Aubert <i>et al.</i>	(BABAR Collab.)
BAUER	04	PR D70 094017	C. Bauer <i>et al.</i>	
HOANG	04	PL B594 127	A.H. Hoang, M. Jamin	
DEDIVITIS	03	NP B675 309	G.M. de Divitis <i>et al.</i>	
EIDEMULLER	03	PR D67 113002	M. Eidemuller	
ERLER	03	PL B558 125	J. Erler, M. Luo	
ZYABLYUK	03	JHEP 0301 081	K.N. Zyblyuk	(ITEP)



$I(J^P) = 0(\frac{1}{2}^+)$

Charge =  $-\frac{1}{3} e$  Bottom =  $-1$

## b-QUARK MASS

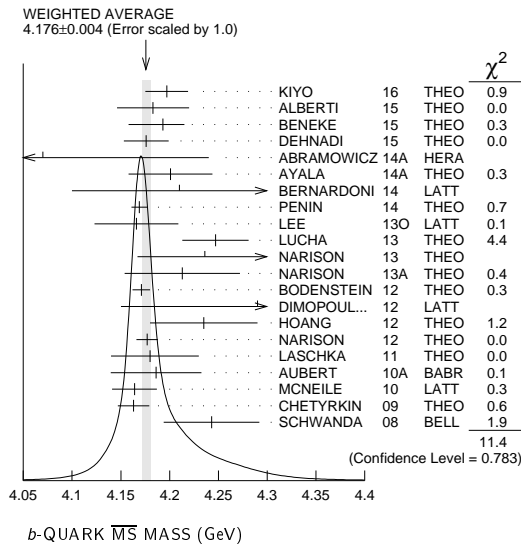
The first value is the “running mass”  $\overline{m}_b(\mu = \overline{m}_b)$  in the  $\overline{MS}$  scheme, and the second value is the  $1S$  mass, which is half the mass of the  $\Upsilon(1S)$  in perturbation theory. For a review of different quark mass definitions and their properties, see EL-KHADRA 02. The  $1S$  mass is better suited for use in analyzing  $B$  decays than the  $\overline{MS}$  mass because it gives a stable perturbative expansion. We have converted masses in other schemes to the  $\overline{MS}$  mass and  $1S$  mass using two-loop QCD perturbation theory with  $\alpha_s(\mu = \overline{m}_b) = 0.223 \pm 0.008$ . The values  $4.18^{+0.04}_{-0.03}$  GeV for the  $\overline{MS}$  mass and  $4.66^{+0.04}_{-0.03}$  GeV for the  $1S$  mass correspond to  $4.78 \pm 0.06$  GeV for the pole mass, using the two-loop conversion formula. A discussion of masses in different schemes can be found in the “Note on Quark Masses.”

$\overline{MS}$ MASS (GeV)	$1S$ MASS (GeV)	DOCUMENT ID	TECN
<b><math>4.18^{+0.04}_{-0.03}</math></b>	<b>OUR EVALUATION</b>	of $\overline{MS}$ Mass. See the ideogram below.	
<b><math>4.66^{+0.04}_{-0.03}</math></b>	<b>OUR EVALUATION</b>	of $1S$ Mass. See the ideogram below.	
$4.197 \pm 0.022$	$4.671 \pm 0.024$	1 KIYO 16	THEO
$4.183 \pm 0.037$	$4.656 \pm 0.041$	2 ALBERTI 15	THEO
$4.193 \pm 0.022$ $-0.035$	$4.667^{+0.024}_{-0.039}$	3 BENEKE 15	THEO
$4.176 \pm 0.023$	$4.648 \pm 0.026$	4 DEHNADI 15	THEO
$4.07 \pm 0.17$	$4.53 \pm 0.19$	5 ABRAMOWICZ14A	HERA
$4.201 \pm 0.043$	$4.676 \pm 0.048$	6 AYALA 14A	THEO
$4.21 \pm 0.11$	$4.69 \pm 0.12$	7 BERNARDONI 14	LATT
$4.169 \pm 0.002 \pm 0.008$	$4.640 \pm 0.002 \pm 0.009$	8 PENIN 14	THEO
$4.166 \pm 0.043$	$4.637 \pm 0.048$	9 LEE 130	LATT
$4.247 \pm 0.034$	$4.727 \pm 0.039$	10 LUCHA 13	THEO
$4.236 \pm 0.069$	$4.715 \pm 0.077$	11 NARISON 13	THEO
$4.213 \pm 0.059$	$4.689 \pm 0.066$	12 NARISON 13A	THEO
$4.171 \pm 0.009$	$4.642 \pm 0.010$	13 BODENSTEIN 12	THEO
$4.29 \pm 0.14$	$4.77 \pm 0.16$	14 DIMOPOUL... 12	LATT
$4.235 \pm 0.003 \pm 0.055$	$4.755 \pm 0.003 \pm 0.058$	15 HOANG 12	THEO
$4.177 \pm 0.011$	$4.649 \pm 0.012$	16 NARISON 12	THEO
$4.18^{+0.05}_{-0.04}$	$4.65^{+0.06}_{-0.04}$	17 LASCHKA 11	THEO
$4.186 \pm 0.044 \pm 0.015$	$4.659 \pm 0.050 \pm 0.017$	18 AUBERT 10A	BABR
$4.164 \pm 0.023$	$4.635 \pm 0.026$	19 MCNEILE 10	LATT
$4.163 \pm 0.016$	$4.633 \pm 0.018$	20 CHETYRKIN 09	THEO
$4.243 \pm 0.049$	$4.723 \pm 0.055$	21 SCHWANDA 08	BELL
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$4.212 \pm 0.032$	$4.688 \pm 0.036$	22 NARISON 12	THEO
$4.171 \pm 0.014$	$4.642 \pm 0.016$	23 NARISON 12A	THEO
$4.173 \pm 0.010$	$4.645 \pm 0.011$	24 NARISON 10	THEO
$5.26 \pm 1.2$	$5.85 \pm 1.3$	25 ABDALLAH 08D	DLPH
$4.42 \pm 0.06 \pm 0.08$	$4.92 \pm 0.07 \pm 0.09$	26 GUAZZINI 08	LATT
$4.347 \pm 0.048 \pm 0.08$	$4.838 \pm 0.053 \pm 0.09$	27 DELLA-MOR... 07	LATT
$4.164 \pm 0.025$	$4.635 \pm 0.028$	28 KUHN 07	THEO
$4.19 \pm 0.40$	$4.66 \pm 0.45$	29 ABDALLAH 06D	DLPH
$4.205 \pm 0.058$	$4.68 \pm 0.06$	30 BOUGHEZAL 06	THEO
$4.20 \pm 0.04$	$4.67 \pm 0.04$	31 BUCHMUEL... 06	THEO

$4.19 \pm 0.06$	$4.66 \pm 0.07$	32 PINEDA 06	THEO
$4.4 \pm 0.3$	$4.9 \pm 0.3$	33,34 GRAY 05	LATT
$4.22 \pm 0.06$	$4.72 \pm 0.07$	35 AUBERT 04x	THEO
$4.17 \pm 0.03$	$4.68 \pm 0.03$	36 BAUER 04	THEO
$4.22 \pm 0.11$	$4.72 \pm 0.12$	34,37 HOANG 04	THEO
$4.25 \pm 0.11$	$4.76 \pm 0.12$	34,38 MCNEILE 04	LATT
$4.22 \pm 0.09$	$4.74 \pm 0.10$	39 BAUER 03	THEO
$4.19 \pm 0.05$	$4.66 \pm 0.05$	40 BORDES 03	THEO
$4.20 \pm 0.09$	$4.67 \pm 0.10$	41 CORCELLA 03	THEO
$4.33 \pm 0.10$	$4.84 \pm 0.11$	34,42 DEDIVITIS 03	LATT
$4.24 \pm 0.10$	$4.72 \pm 0.11$	43 EIDEMULLER 03	THEO
$4.207 \pm 0.031$	$4.682 \pm 0.035$	44 ERLER 03	THEO
$4.33 \pm 0.06 \pm 0.10$	$4.82 \pm 0.07 \pm 0.11$	45 MAHMOOD 03	CLEO
$4.190 \pm 0.032$	$4.663 \pm 0.036$	46 BRAMBILLA 02	THEO
$4.346 \pm 0.070$	$4.837 \pm 0.078$	47 PENIN 02	THEO

- KIYO 16 determine  $\overline{m}_b(\overline{m}_b)$  from the  $\Upsilon(1S)$  mass at order  $\alpha_s^3$  (N3LO). We have converted this to the  $1S$  scheme.
- ALBERTI 15 determine  $\overline{m}_b(\overline{m}_b)$  from fits to inclusive  $B \rightarrow X_c e \overline{\nu}$  decay. We have converted this to the  $1S$  scheme. They also find  $m_b^{\text{kin}}(1 \text{ GeV}) = 4.553 \pm 0.020 \text{ GeV}$ .
- BENEKE 15 determine  $\overline{m}_b(\overline{m}_b)$  using sum rules for  $e^+e^- \rightarrow \text{hadrons}$  at order N3LO including finite  $m_c$  effects. We have converted this to the  $1S$  scheme. They also find  $m_b^{\text{PS}}(2 \text{ GeV}) = 4.532^{+0.013}_{-0.039} \text{ GeV}$ . When the four-loop conversion between the pole and the  $\overline{MS}$  mass is applied in BENEKE 16, the  $\overline{m}_b(\overline{m}_b)$  mass changes to  $4.203^{+0.016}_{-0.034} \text{ GeV}$ .
- DEHNADI 15 determine  $\overline{m}_b(\overline{m}_b)$  using sum rules for  $e^+e^- \rightarrow \text{hadrons}$  at order  $\alpha_s^3$  (N3LO), and fitting to both experimental data and lattice results. We have converted this to the  $1S$  scheme.
- ABRAMOWICZ 14A determine  $\overline{m}_b(\overline{m}_b) = 4.07 \pm 0.14^{+0.01+0.05+0.08}_{-0.07-0.00-0.05}$  from the production of  $b$  quarks in  $ep$  collisions at HERA. The errors due to fitting, modeling, PDF parameterization, and theoretical QCD uncertainties due to the values of  $\alpha_s$ ,  $m_c$ , and the renormalization scale  $\mu$  have been combined in quadrature. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- AYALA 14A determine  $\overline{m}_b(\overline{m}_b)$  from the  $\Upsilon(1S)$  mass computed to N<sup>3</sup>LO order in perturbation theory using a renormalon subtracted scheme. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- BERNARDONI 14 determine  $m_b$  from  $N_f = 2$  lattice calculations using heavy quark effective theory non-perturbatively renormalized and matched to QCD at  $1/m$  order. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- PENIN 14 determine  $\overline{m}_b(\overline{m}_b) = 4.169 \pm 0.008 \pm 0.002 \pm 0.002$  using an estimate of the order  $\alpha_s^3$   $b$ -quark vacuum polarization function in the threshold region, including finite  $m_c$  effects. The errors of  $\pm 0.008$  from theoretical uncertainties, and  $\pm 0.002$  from  $\alpha_s$  have been combined in quadrature. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- LEE 130 determines  $m_b$  using lattice calculations of the  $\Upsilon$  and  $B_s$  binding energies in NRQCD, including three light dynamical quark flavors. The quark mass shift in NRQCD is determined to order  $\alpha_s^2$ , with partial  $\alpha_s^3$  contributions.
- LUCHA 13 determines  $m_b$  from QCD sum rules for heavy-light currents using the lattice value for  $\Gamma_B$  of  $191.5 \pm 7.3 \text{ GeV}$ .
- NARISON 13 determines  $m_b$  using QCD spectral sum rules to order  $\alpha_s^2$  (NNLO) and including condensates up to dimension 6. We have converted the  $\overline{MS}$  value to the  $1S$  scheme.
- NARISON 13A determines  $m_b$  using HQET sum rules to order  $\alpha_s^2$  (NNLO) and the  $B$  meson mass and decay constant.
- BODENSTEIN 12 determine  $m_b$  using sum rules for the vector current correlator and the  $e^+e^- \rightarrow Q\overline{Q}$  total cross-section. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- DIMOPOULOS 12 determine quark masses from a lattice computation using  $N_f = 2$  dynamical flavors of twisted mass fermions. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- HOANG 12 determine  $m_b$  using non-relativistic sum rules for the  $\Upsilon$  system at order  $\alpha_s^2$  (NNLO) with renormalization group improvement.
- Determines  $m_b$  to order  $\alpha_s^3$  (N3LO), including the effect of gluon condensates up to dimension eight combining the methods of NARISON 12 and NARISON 12A. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- LASCHKA 11 determine the  $b$  mass from the charmonium spectrum. The theoretical computation uses the heavy  $Q\overline{Q}$  potential to order  $1/m_Q$  obtained by matching the short-distance perturbative result onto lattice QCD result at larger scales. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- AUBERT 10A determine the  $b$ - and  $c$ -quark masses from a fit to the inclusive decay spectra in semileptonic  $B$  decays in the kinetic scheme (and convert it to the  $\overline{MS}$  scheme). We have converted this to the  $1S$  scheme.
- MCNEILE 10 determines  $m_b$  by comparing order  $\alpha_s^3$  (N3LO) perturbative results for the pseudo-scalar current to lattice simulations with  $N_f = 2+1$  sea-quarks by the HPQCD collaboration. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- CHETYRKIN 09 determine  $m_c$  and  $m_b$  from the  $e^+e^- \rightarrow Q\overline{Q}$  cross-section and sum rules, using an order  $\alpha_s^3$  (N3LO) computation of the heavy quark vacuum polarization. We have converted their  $m_b$  to the  $1S$  scheme.
- SCHWANDA 08 measure moments of the inclusive photon spectrum in  $B \rightarrow X_s \gamma$  decay to determine  $m_b^{1S}$ . We have converted this to  $\overline{MS}$  scheme.
- NARISON 12 determines  $m_b$  using exponential sum rules for the vector current correlator to order  $\alpha_s^3$ , including the effect of gluon condensates up to dimension eight. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- NARISON 12A determines  $m_b$  using sum rules for the vector current correlator to order  $\alpha_s^3$ , including the effect of gluon condensates up to dimension eight. We have converted  $\overline{m}_b(\overline{m}_b)$  to the  $1S$  scheme.
- NARISON 10 determines  $m_b$  from ratios of moments of vector current correlators computed to order  $\alpha_s^3$  and including the dimension-six gluon condensate. These values are taken from the erratum to that reference.

- 25 ABDALLAH 08d determine  $\overline{m}_b(M_Z) = 3.76 \pm 1.0$  GeV from a leading order study of four-jet rates at LEP. We have converted this to  $\overline{m}_b(\overline{m}_b)$  and  $m_b^{1S}$ .
- 26 GUAZZINI 08 determine  $\overline{m}_b(\overline{m}_b)$  from a quenched lattice simulation of heavy meson masses. The  $\pm 0.08$  is an estimate of the quenching error. We have converted these values to the 1S scheme.
- 27 DELLA-MORTE 07 determine  $\overline{m}_b(\overline{m}_b)$  from a computation of the spin-averaged  $B$  meson mass using quenched lattice HQET at order  $1/m$ . The  $\pm 0.08$  is an estimate of the quenching error.
- 28 KUHN 07 determine  $\overline{m}_b(\mu = 10 \text{ GeV}) = 3.609 \pm 0.025$  GeV and  $\overline{m}_b(\overline{m}_b)$  from a four-loop sum-rule computation of the cross-section for  $e^+e^- \rightarrow$  hadrons in the bottom threshold region. We have converted this to the 1S scheme.
- 29 ABDALLAH 06d determine  $m_b(M_Z) = 2.85 \pm 0.32$  GeV from  $Z$ -decay three-jet events containing a  $b$ -quark. We have converted this to  $\overline{m}_b(\overline{m}_b)$  and  $m_b^{1S}$ .
- 30 BOUGHEZAL 06  $\overline{MS}$  scheme result comes from the first moment of the hadronic production cross-section to order  $\alpha_s^3$ . We have converted it to the 1S scheme.
- 31 BUCHMUELLER 06 determine  $m_b$  and  $m_c$  by a global fit to inclusive  $B$  decay spectra. We have converted this to the 1S scheme.
- 32 PINEDA 06  $\overline{MS}$  scheme result comes from a partial NNLL evaluation (complete at order  $\alpha_s^2$  (NNLO)) of sum rules of the bottom production cross-section in  $e^+e^-$  annihilation. We have converted it to the 1S scheme.
- 33 GRAY 05 determines  $\overline{m}_b(\overline{m}_b)$  from a lattice computation of the  $\Upsilon$  spectrum. The simulations have 2+1 dynamical light flavors. The  $b$  quark is implemented using NRQCD. We have converted  $m_b$  to the 1S scheme.
- 34 AUBERT 04x obtain  $m_b$  from a fit to the hadron mass and lepton energy distributions in semileptonic  $B$  decay. The paper quotes values in the kinetic scheme. The  $\overline{MS}$  value has been provided by the BABAR collaboration, and we have converted this to the 1S scheme.
- 35 BAUER 04 determine  $m_b$ ,  $m_c$  and  $m_b - m_c$  by a global fit to inclusive  $B$  decay spectra.
- 36 HOANG 04 determines  $\overline{m}_b(\overline{m}_b)$  from moments at order  $\alpha_s^2$  of the bottom production cross-section in  $e^+e^-$  annihilation.
- 37 MCNEILE 04 use lattice QCD with dynamical light quarks and a static heavy quark to compute the masses of heavy-light mesons.
- 38 BAUER 03 determine the  $b$  quark mass by a global fit to  $B$  decay observables. The experimental data includes lepton energy and hadron invariant mass moments in semileptonic  $B \rightarrow X_c \ell \nu_\ell$  decay, and the inclusive photon spectrum in  $B \rightarrow X_s \gamma$  decay. The theoretical expressions used are of order  $1/m^3$ , and  $\alpha_s^2 \beta_0$ .
- 40 BORDES 03 determines  $m_b$  using QCD finite energy sum rules to order  $\alpha_s^2$ .
- 41 CORCELLA 03 determines  $\overline{m}_b$  using sum rules computed to order  $\alpha_s^2$ . Includes charm quark mass effects.
- 42 DEDIVITIIS 03 use a quenched lattice computation of heavy-heavy and heavy-light meson masses.
- 43 EIDEMULLER 03 determines  $\overline{m}_b$  and  $\overline{m}_c$  using QCD sum rules.
- 44 ERLER 03 determines  $\overline{m}_b$  and  $\overline{m}_c$  using QCD sum rules. Includes recent BES data.
- 45 MAHMOOD 03 determines  $m_b^{1S}$  by a fit to the lepton energy moments in  $B \rightarrow X_c \ell \nu_\ell$  decay. The theoretical expressions used are of order  $1/m^3$  and  $\alpha_s^2 \beta_0$ . We have converted their result to the  $\overline{MS}$  scheme.
- 46 BRAMBILLA 02 determine  $\overline{m}_b(\overline{m}_b)$  from a computation of the  $\Upsilon(1S)$  mass to order  $\alpha_s^4$ , including finite  $m_c$  corrections. We have converted this to the 1S scheme.
- 47 PENIN 02 determines  $\overline{m}_b$  from the spectrum of the  $\Upsilon$  system.



## b-QUARK REFERENCES

BENEKE	16	arXiv:1601.02949	M. Beneke <i>et al.</i>
KIYO	16	PL B752 122	Y. Kiyo, G. Mishima, Y. Sumino
ALBERTI	15	PRL 114 061802	A. Alberti <i>et al.</i>
BENEKE	15	NP B891 42	M. Beneke <i>et al.</i>
DEHNADI	15	JHEP 1508 155	B. Dehnadi, A.H. Hoang, V. Mateu
ABRAMOWICZ	14A	JHEP 1409 127	H. Abramowicz <i>et al.</i> (ZEUS Collab.)
AYALA	14A	JHEP 1409 045	C. Ayala, G. Cvetik, A. Pineda
BERNARDONI	14	PL B730 171	F. Bernardoni <i>et al.</i> (ALPHA Collab.)
PENIN	14	JHEP 1404 120	A.A. Penin, N. Zerf

LEE	13O	PR D87 074018	A.J. Lee <i>et al.</i> (HPQCD Collab.)
LUCHA	13	PR D88 056011	W. Lucha, D. Melikhov, S. Simula (WIEN, MOSU+)
NARISON	13	PL B718 1321	S. Narison (MONP)
NARISON	13A	PL B721 269	S. Narison (MONP)
BODENSTEIN	12	PR D85 034003	S. Bodenstein <i>et al.</i> (CAPE, VALE, MANZ+)
DIMOPOULOS	12	JHEP 1201 046	P. Dimopoulos <i>et al.</i> (ETM Collab.)
HOANG	12	JHEP 1210 188	A.H. Hoang, P. Ruiz-Femenia, M. Stahlhofen (WIEN+)
NARISON	12	PL B707 259	S. Narison (MONP)
NARISON	12A	PL B706 412	S. Narison (MONP)
LASCHKA	11	PR D83 094002	A. Laschka, N. Kaiser, W. Weise
AUBERT	10A	PR D81 032003	B. Aubert <i>et al.</i> (BABAR Collab.)
MCNEILE	10	PR D82 034512	C. McNeile <i>et al.</i> (HPQCD Collab.)
NARISON	10	PL B693 559	S. Narison (MONP)
Also		PL B705 544 (errat.)	S. Narison (MONP)
CHETYRKIN	09	PR D80 074010	K.G. Chetyrkin <i>et al.</i> (KARL, BNL)
ABDALLAH	08D	EPJ C55 525	J. Abdallah <i>et al.</i> (DELPHI Collab.)
GUAZZINI	08	JHEP 0801 076	D. Guazzini, R. Sommer, N. Tantalo
SCHWANDA	08	PR D78 032016	C. Schwanda <i>et al.</i> (BELLE Collab.)
DELLA-MORTE	07	JHEP 0701 007	M. Della Morte <i>et al.</i>
KUHN	07	NP B778 192	J.H. Kuhn, M. Steinhauser, C. Sturm
ABDALLAH	06D	EPJ C46 569	J. Abdallah <i>et al.</i> (DELPHI Collab.)
BOUGHEZAL	06	PR D74 074006	R. Boughezal, M. Czakon, T. Schutzmeier
BUCHMUELLER	06	PR D73 073008	O.L. Buchmuller, H.U. Flacher (RHBL)
PINEDA	06	PR D73 111501	A. Pineda, A. Signer
GRAY	05	PR D72 094507	A. Gray <i>et al.</i> (HPQCD, UKQCD Collab.)
AUBERT	04X	PRL 93 011803	B. Aubert <i>et al.</i> (BABAR Collab.)
BAUER	04	PR D70 094017	C. Bauer <i>et al.</i>
HOANG	04	PL B594 127	A.H. Hoang, M. Jamin
MCNEILE	04	PL B600 77	C. McNeile, C. Michael, G. Thompson (UKQCD Collab.)
BAUER	03	PR D67 054012	C.W. Bauer <i>et al.</i>
BORDES	03	PL B562 81	J. Bordes, J. Penarrocha, K. Schilcher
CORCELLA	03	PL B554 133	G. Corcella, A.H. Hoang
DEDIVITIIS	03	NP B675 309	G.M. de Divitiis <i>et al.</i>
EIDEMULLER	03	PR D67 113002	M. Eidemuller
ERLER	03	PL B558 125	J. Erler, M. Luo
MAHMOOD	03	PR D67 072001	A.H. Mahmood <i>et al.</i> (CLEO Collab.)
BRAMBILLA	02	PR D65 034001	N. Brambilla, Y. Sumino, A. Vairo
EL-KHADRA	02	ARNPS 52 201	A.X. El-Khadra, M. Luke
PENIN	02	PL B538 335	A. Penin, M. Steinhauser



$$I(J^P) = 0(\frac{1}{2}^{+})$$

$$\text{Charge} = \frac{2}{3} e \quad \text{Top} = +1$$

## THE TOP QUARK

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## A. Introduction

The top quark is the  $Q = 2/3$ ,  $T_3 = +1/2$  member of the weak-isospin doublet containing the bottom quark (see the review on the “Electroweak Model and Constraints on New Physics” for more information). Its phenomenology is driven by its large mass. Being heavier than a  $W$  boson, it is the only quark that decays semi-weakly, i.e., into a real  $W$  boson and a  $b$  quark. Therefore, it has a very short lifetime and decays before hadronization can occur. In addition, it is the only quark whose Yukawa coupling to the Higgs boson is order of unity. For these reasons the top quark plays a special role in the Standard Model (SM) and in many extensions thereof. Its phenomenology provides a unique laboratory where our understanding of the strong interactions, both in the perturbative and non-perturbative regimes, can be tested. An accurate knowledge of its properties (mass, couplings, production cross section, decay branching ratios, *etc.*) can bring key information on fundamental interactions at the electroweak breaking scale and beyond. This review provides a concise discussion of the experimental and theoretical issues involved in the determination the top-quark properties.

## B. Top-quark production at the Tevatron and LHC

In hadron collisions, top quarks are produced dominantly in pairs through the processes  $q\bar{q} \rightarrow t\bar{t}$  and  $gg \rightarrow t\bar{t}$ , at leading order in QCD. Approximately 85% of the production cross section at the Tevatron is from  $q\bar{q}$  annihilation, with

# Quark Particle Listings

*t*

the remainder from gluon-gluon fusion, while at LHC energies about 90% of the production is from the latter process at  $\sqrt{s} = 14$  TeV ( $\approx 80\%$  at  $\sqrt{s} = 7$  TeV).

Predictions for the total cross sections are now available at next-to-next-to leading order (NNLO) with next-to-next-to-leading-log (NNLL) soft gluon resummation [1]. These results supersede previous approximate ones [2]. Assuming a top-quark mass of 173.3 GeV/c<sup>2</sup>, close to the Tevatron + LHC average [3] (LHC results not yet included), the resulting theoretical prediction of the top-quark pair cross-section at NNLO+NNLL accuracy at the Tevatron at  $\sqrt{s} = 1.96$  TeV is  $\sigma_{t\bar{t}} = 7.16^{+0.11+0.17}_{-0.20-0.12}$  pb where the first uncertainty is from scale dependence and the second from parton distribution functions. At the LHC, assuming a top-quark mass of 173.2 GeV/c<sup>2</sup> the cross sections are:  $\sigma_{t\bar{t}} = 173.6^{+4.5+8.9}_{-5.9-8.9}$  pb, at  $\sqrt{s} = 7$  TeV,  $\sigma_{t\bar{t}} = 247.7^{+6.3+11.5}_{-8.5-11.5}$  pb at  $\sqrt{s} = 8$  TeV, and  $\sigma_{t\bar{t}} = 816.0^{+19.4+34.4}_{-28.6-34.4}$  pb at  $\sqrt{s} = 13$  TeV [1].

Electroweak single top-quark production mechanisms, namely from  $q\bar{q}' \rightarrow t\bar{b}$  [4],  $qb \rightarrow q't$  [5], mediated by virtual *s*-channel and *t*-channel *W*-bosons, and *Wt*-associated production, through  $bg \rightarrow W^-t$ , lead to somewhat smaller cross sections. For example, *t*-channel production, while suppressed by the weak coupling with respect to the strong pair production, is kinematically enhanced, resulting in a sizable cross section both at Tevatron and LHC energies. At the Tevatron, the *t*- and *s*-channel cross sections of top and antitop are identical, while at the LHC they are not, due to the charge-asymmetric initial state. Approximate NNLO cross sections for *t*-channel single top-quark production ( $t + \bar{t}$ ) are calculated for  $m_t = 173.3$  GeV/c<sup>2</sup> to be  $2.06^{+0.13}_{-0.13}$  pb in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV (scale and parton distribution functions uncertainties are combined in quadrature) and  $65.7^{+1.9}_{-1.9}$  ( $87.1^{+0.24}_{-0.24}$ ) pb in  $pp$  collisions at  $\sqrt{s} = 7$  (8) TeV, where 65% and 35% are the relative proportions of *t* and  $\bar{t}$  [6]. A calculation at NNLO accuracy for the *t*-channel cross section has been recently performed predicting a cross section of  $85.1^{+2.5}_{-1.4}$  pb at 8 TeV [7]. For the *s*-channel, these calculations yield  $1.03^{+0.05}_{-0.05}$  pb for the Tevatron, and  $4.5^{+0.2}_{-0.2}$  ( $5.5^{+0.2}_{-0.2}$ ) pb for  $\sqrt{s} = 7$  (8) TeV at the LHC, with 69% (31%) of top (anti-top) quarks [8]. While negligible at the Tevatron, at LHC energies the *Wt*-associated production becomes relevant. At  $\sqrt{s} = 7$  (8) TeV, an approximate NNLO calculation gives  $15.5^{+1.2}_{-1.2}$  ( $22.1^{+1.5}_{-1.5}$ ) pb ( $t + \bar{t}$ ), with an equal proportion of top and anti-top quarks [9].

Assuming  $|V_{tb}| \gg |V_{td}|, |V_{ts}|$  (see the review “The CKM Quark-Mixing Matrix” for more information), the cross sections for single top production are proportional to  $|V_{tb}|^2$ , and no extra hypothesis is needed on the number of quark families or on the unitarity of the CKM matrix in extracting  $|V_{tb}|$ . Separate measurements of the *s*- and *t*-channel processes provide sensitivity to physics beyond the Standard Model [10].

With a mass above the *Wb* threshold, and  $|V_{tb}| \gg |V_{td}|, |V_{ts}|$ , the decay width of the top quark is expected to be dominated by the two-body channel  $t \rightarrow Wb$ . Neglecting terms

of order  $m_b^2/m_t^2$ ,  $\alpha_s^2$ , and  $(\alpha_s/\pi)M_W^2/m_t^2$ , the width predicted in the SM at NLO is [11]:

$$\Gamma_t = \frac{G_F m_t^3}{8\pi\sqrt{2}} \left(1 - \frac{M_W^2}{m_t^2}\right)^2 \left(1 + 2\frac{M_W^2}{m_t^2}\right) \left[1 - \frac{2\alpha_s}{3\pi} \left(\frac{2\pi^2}{3} - \frac{5}{2}\right)\right], \quad (1)$$

where  $m_t$  refers to the top-quark pole mass. The width for a value of  $m_t = 173.3$  GeV/c<sup>2</sup> is 1.35 GeV/c<sup>2</sup> (we use  $\alpha_s(M_Z) = 0.118$ ) and increases with mass. With its correspondingly short lifetime of  $\approx 0.5 \times 10^{-24}$  s, the top quark is expected to decay before top-flavored hadrons or  $t\bar{t}$ -quarkonium-bound states can form [12]. In fact, since the decay time is close to the would-be-resonance binding time, a peak will be visible in  $e^+e^-$  scattering at the  $t\bar{t}$  threshold [13] and it is in principle present (yet very difficult to measure) in hadron collisions, too [14]. The order  $\alpha_s^2$  QCD corrections to  $\Gamma_t$  are also available [15], thereby improving the overall theoretical accuracy to better than 1%.

The final states for the leading pair-production process can be divided into three classes:

- A.  $t\bar{t} \rightarrow W^+ b W^- \bar{b} \rightarrow q\bar{q}' b q'' \bar{q}'' \bar{b}$ , (45.7%)
- B.  $t\bar{t} \rightarrow W^+ b W^- \bar{b} \rightarrow q\bar{q}' b \ell^- \bar{\nu}_\ell \bar{b} + \ell^+ \nu_\ell b q'' \bar{q}'' \bar{b}$ , (43.8%)
- C.  $t\bar{t} \rightarrow W^+ b W^- \bar{b} \rightarrow \ell^+ \nu_\ell b \ell'^- \bar{\nu}_{\ell'}$ . (10.5%)

The quarks in the final state evolve into jets of hadrons. A, B, and C are referred to as the all-jets, lepton+jets ( $\ell$ +jets), and dilepton ( $\ell\ell$ ) channels, respectively. Their relative contributions, including hadronic corrections, are given in parentheses assuming lepton universality. While  $\ell$  in the above processes refers to *e*,  $\mu$ , or  $\tau$ , most of the analyses distinguish the *e* and  $\mu$  from the  $\tau$  channel, which is more difficult to reconstruct. Therefore, in what follows, we will use  $\ell$  to refer to *e* or  $\mu$ , unless otherwise noted. Here, typically leptonic decays of  $\tau$  are included. In addition to the quarks resulting from the top-quark decays, extra QCD radiation (quarks and gluons) from the colored particles in the event can lead to extra jets.

The number of jets reconstructed in the detectors depends on the decay kinematics, as well as on the algorithm for reconstructing jets used by the analysis. Information on the transverse momenta of neutrinos is obtained from the imbalance in transverse momentum measured in each event (missing  $p_T$ , which is here also called missing  $E_T$ ).

The identification of top quarks in the electroweak single top channel is much more difficult than in the QCD  $t\bar{t}$  channel, due to a less distinctive signature and significantly larger backgrounds, mostly due to  $t\bar{t}$  and *W*+jets production.

Fully exclusive predictions via Monte Carlo generators for the  $t\bar{t}$  and single top production processes at NLO accuracy in QCD, including top-quark decays, are available [16,17] through the MC@NLO [18] and POWHEG [19] methods.

Besides fully inclusive QCD or EW top-quark production, more exclusive final states can be accessed at hadron colliders, whose cross sections are typically much smaller, yet can provide key information on the properties of the top quark. For all relevant final states (*e.g.*,  $t\bar{t}V$ ,  $t\bar{t}VV$  with  $V = \gamma, W, Z$ ,  $t\bar{t}H$ ,  $t\bar{t}$ +jets,  $t\bar{t}b\bar{b}$ ,  $t\bar{t}t\bar{t}$ ) automatic or semi-automatic predictions



at NLO accuracy in QCD also in the form of event generators, *i.e.*, interfaced to parton-shower programs, are available (see the review “Monte Carlo event generators” for more information).

### C. Top-quark measurements

Since the discovery of the top quark, direct measurements of  $t\bar{t}$  production have been made at five center-of-mass energies, providing stringent tests of QCD. The first measurements were made in Run I at the Tevatron at  $\sqrt{s} = 1.8$  TeV. In Run II at the Tevatron relatively precise measurements were made at  $\sqrt{s} = 1.96$  TeV. Finally, beginning in 2010, measurements have been made at the LHC at  $\sqrt{s} = 7$  TeV and  $\sqrt{s} = 8$  TeV, and very recently at  $\sqrt{s} = 13$  TeV.

Production of single top quarks through electroweak interactions has now been measured with good precision at the Tevatron at  $\sqrt{s} = 1.96$  TeV, and at the LHC at  $\sqrt{s} = 7$  TeV and  $\sqrt{s} = 8$  TeV, and now also at  $\sqrt{s} = 13$  TeV. Recent measurements at the Tevatron have managed to separate the *s*- and *t*-channel production cross sections, and at the LHC, the *Wt* mechanism as well, though the *t*-channel is measured with best precision to date. The measurements allow an extraction of the CKM matrix element  $V_{tb}$ .

With approximately  $10 \text{ fb}^{-1}$  of Tevatron data analyzed as of this writing, and almost  $5 \text{ fb}^{-1}$  at 7 TeV,  $20 \text{ fb}^{-1}$  at 8 TeV and the first  $78 \text{ pb}^{-1}$  at 13 TeV at the LHC, many properties of the top quark have been measured with precision. These include properties related to the production mechanism, such as  $t\bar{t}$  spin correlations, forward-backward or charge asymmetries, and differential production cross sections, as well as properties related to the  $tWb$  decay vertex, such as the helicity of the *W*-bosons from the top-quark decay. Recently, also studies of the  $t\bar{t}\gamma$  and the  $t\bar{t}Z$  interactions have been made. In addition, many searches for physics beyond the Standard Model are being performed with increasing reach in both production and decay channels.

In the following sections we review the current status of measurements of the characteristics of the top quark.

### C.1 Top-quark production

**C.1.1  $t\bar{t}$  production:** Fig. 1 summarizes the  $t\bar{t}$  production cross-section measurements from both the Tevatron and LHC. The most recent measurement from DØ [20], combining the measurements from the dilepton and lepton plus jets final states in  $9.7 \text{ fb}^{-1}$ , is  $7.73 \pm 0.13 \pm 0.55 \text{ pb}$ .

From CDF the most precise measurement made recently [21] is in  $8.8 \text{ fb}^{-1}$  in the dilepton channel requiring at least one *b*-tag, yielding  $7.09 \pm 0.84 \text{ pb}$ . Both of these measurements assume a top-quark mass of  $172.5 \text{ GeV}/c^2$ . The dependence of the cross section measurements on the value chosen for the mass is less than that of the theory calculations because it only affects the determination of the acceptance. In some analyses also the shape of topological variables might be modified.

The resulting combined  $t\bar{t}$  cross-section is  $\sigma_{t\bar{t}} = 7.63 \pm 0.50 \text{ pb}$  (6.6%) for CDF,  $\sigma_{t\bar{t}} = 7.56 \pm 0.59 \text{ pb}$  (7.8%) for DØ and

$\sigma_{t\bar{t}} = 7.60 \pm 0.41 \text{ pb}$  (5.4%) for the Tevatron combination [22] in good agreement with the SM expectation of  $7.35^{+0.28}_{-0.33} \text{ pb}$  at NNLO+NNLL in perturbative QCD [1] for a top mass of  $172.5 \text{ GeV}$ . The contributions to the uncertainty are  $0.20 \text{ pb}$  from statistical sources,  $0.29 \text{ pb}$  from systematic sources, and  $0.21 \text{ pb}$  from the uncertainty on the integrated luminosity.

CDF has measured the  $t\bar{t}$  production cross section in the dilepton channel with one hadronically decaying tau in  $9.0 \text{ fb}^{-1}$ , yielding  $\sigma_{t\bar{t}} = 8.1 \pm 2.1 \text{ pb}$ . By separately identifying the single-tau and the ditau components, they measure the branching fraction of the top quark into the tau lepton, tau neutrino, and bottom quark to be  $(9.6 \pm 2.8)\%$  [23]. CDF also performs measurements of the  $t\bar{t}$  production cross section normalized to the *Z* production cross section in order to reduce the impact of the luminosity uncertainty.

The LHC experiments ATLAS and CMS use similar techniques to measure the  $t\bar{t}$  cross-section in *pp* collisions. The most precise measurements come from the dilepton channel, and in particular the  $e\mu$  channel. At  $\sqrt{s} = 7$  TeV, ATLAS uses  $4.6 \text{ fb}^{-1}$  of  $e\mu$  events in which they select an extremely clean sample and determine the  $t\bar{t}$  cross-section simultaneously with the efficiency to reconstruct and tag *b*-jets, yielding  $\sigma_{t\bar{t}} = 182.9 \pm 7.1 \text{ pb}$ , corresponding to 3.9% precision [24]. Other measurements by ATLAS at  $\sqrt{s} = 7$  TeV, include a measurement in  $0.7 \text{ fb}^{-1}$  in the lepton+jets channel [25], in the dilepton channel [26], and in  $1.02 \text{ fb}^{-1}$  in the all-hadronic channel [27], which together yield a combined value of  $\sigma_{t\bar{t}} = 177 \pm 3(\text{stat.})^{+8}_{-7}(\text{syst.}) \pm 7(\text{lumi.}) \text{ pb}$  (6.2%) assuming  $m_t = 172.5 \text{ GeV}/c^2$  [28]. In  $4.7 \text{ fb}^{-1}$  of all-jets events, they obtain  $\sigma_{t\bar{t}} = 168 \pm 62 \text{ pb}$  [29]. Further analyses in the hadronic  $\tau$  plus jets channel in  $1.67 \text{ fb}^{-1}$  [30] and the hadronic  $\tau + \text{lepton}$  channel in  $2.05 \text{ fb}^{-1}$  [31], yield consistent albeit less precise results. The most precise measurement from CMS is also obtained in the dilepton channel, where they measure  $\sigma_{t\bar{t}} = 162 \pm 2(\text{stat.}) \pm 5(\text{syst.}) \pm 4(\text{lumi.}) \text{ pb}$ , corresponding to a 4.2% precision [32]. Other measurements at  $\sqrt{s} = 7$  TeV from CMS include measurements with  $2.3 \text{ fb}^{-1}$  in the  $e/\mu$ +jets channel [33], with  $3.5 \text{ fb}^{-1}$  in the all-hadronic channel [34], with  $2.2 \text{ fb}^{-1}$  in the lepton+ $\tau$  channel [35], and with  $3.9 \text{ fb}^{-1}$  in the  $\tau$ +jets channel [36]. ATLAS and CMS also provide a combined cross section of  $173.3 \pm 2.3(\text{stat.}) \pm 7.6(\text{syst.}) \pm 6.3(\text{lump.}) \text{ pb}$  using slightly older results based on  $0.7 - 1.1 \text{ fb}^{-1}$  [37].

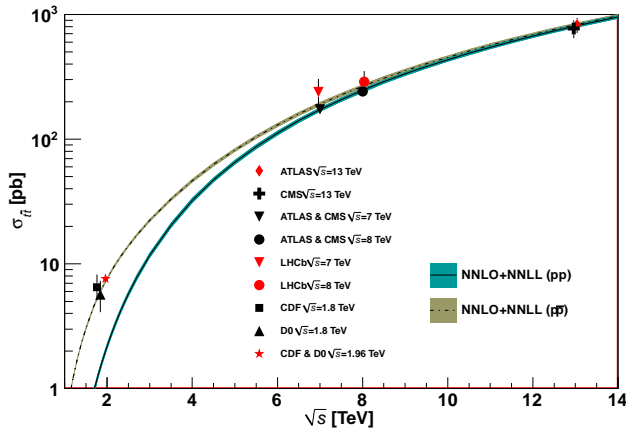
At  $\sqrt{s} = 8$  TeV, ATLAS measures the  $t\bar{t}$  cross-section with  $20.3 \text{ fb}^{-1}$  using  $e\mu$  dilepton events, with a simultaneous measurement of the *b*-tagging efficiency, yielding  $\sigma_{t\bar{t}} = 242.4 \pm 1.7(\text{stat.}) \pm 5.5(\text{syst.}) \pm 7.5(\text{lumi.}) \pm 4.2(\text{beam energy}) \text{ pb}$  [24] assuming  $m_t = 172.5 \text{ GeV}/c^2$ , which corresponds to a 4.7% precision. In the lepton+jets channel, they measure  $\sigma_{t\bar{t}} = 260 \pm 1(\text{stat.})^{+20}_{-23}(\text{syst.}) \pm 8(\text{lumi.}) \pm 4(\text{beam energy}) \text{ pb}$  [38] in  $20.3 \text{ fb}^{-1}$  using a likelihood discriminant fit and *b*-jet identification.

CMS performs a template fit to the  $M_{b\bar{b}}$  mass distribution using  $2.8 \text{ fb}^{-1}$  in the lepton+jets channel yielding

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$\sigma_{t\bar{t}} = 228 \pm 9(\text{stat.})^{+29}_{-26}(\text{syst.}) \pm 10(\text{lumi.})$  pb [39]. In the dilepton channel, the cross sections are extracted using a binned likelihood fit to multi-differential final state distributions related to identified  $b$  quark and other jets in the event. Using the full data samples collected in 2011 and 2012 they obtain  $\sigma_{t\bar{t}} = 245.6 \pm 1.3(\text{stat.}) \pm 6.0(\text{syst.}) \pm 6.5(\text{lumi.})$  pb [40]. The cross section is also measured in the all-jets final state giving  $\sigma_{t\bar{t}} = 275.6 \pm 6.1(\text{stat.}) \pm 37.8(\text{syst.}) \pm 7.2(\text{lumi.})$  pb [41]. In combination of the most precise  $e\mu$  measurements in  $5.3 - 20.3 \text{ fb}^{-1}$ , ATLAS and CMS together yield  $\sigma_{t\bar{t}} = 241.5 \pm 1.4(\text{stat.}) \pm 5.7(\text{syst.}) \pm 6.2(\text{lumi.})$  pb [42], which corresponds to a 3.5% precision, challenging the precision of the corresponding theoretical predictions.



**Figure 1:** Measured and predicted  $t\bar{t}$  production cross sections from Tevatron energies in  $p\bar{p}$  collisions to LHC energies in  $pp$  collisions. Tevatron data points at  $\sqrt{s} = 1.8$  TeV are from Refs. [49,50]. Those at  $\sqrt{s} = 1.96$  TeV are from Refs. [20–22]. The ATLAS, CMS, and LHCb data points are from Refs. [28–29,38], and [33–34], and [43], respectively. Theory curves and uncertainties are generated using [1] for  $m_t = 172.5 \text{ GeV}/c^2$ , the  $m_t$  value assumed in the cross section measurements. Figure adapted from Ref. [46].

Recently, the LHCb collaboration presented the first observation of top-quark production in the forward region in  $pp$ -collisions. The  $W + b$  final state with  $W \rightarrow \mu\nu$  is reconstructed using muons with a transverse momentum,  $p_T$ , larger than 25 GeV in the pseudorapidity range  $2.0 < \eta < 4.5$ . The  $b$ -jets are required to have  $50 \text{ GeV} < p_T < 100 \text{ GeV}$  and  $2.2 < \eta < 4.2$ , while the transverse component of the sum of the muon and  $b$ -jet momenta must satisfy  $p_T > 20 \text{ GeV}$ . The results are based on data corresponding to integrated luminosities of 1.0 and  $2.0 \text{ fb}^{-1}$  collected at center-of-mass energies of 7 and 8 TeV by LHCb. The inclusive top quark production cross-sections in the fiducial region are  $\sigma_{t\bar{t}} = 239 \pm 53(\text{stat.}) \pm 38(\text{syst.})$  pb at 7 TeV, and  $\sigma_{t\bar{t}} = 289 \pm 43(\text{stat.}) \pm 46(\text{syst.})$  pb at 8 TeV [43].

Very recently, ATLAS and CMS have also measured the  $t\bar{t}$  production cross section with early Run-II data at  $\sqrt{s} = 13 \text{ TeV}$  in  $e\mu$  events with at least one  $b$ -tag. ATLAS uses  $78 \text{ pb}^{-1}$

and obtains  $\sigma_{t\bar{t}} = 825 \pm 114$  pb [44]. CMS uses  $42 \text{ pb}^{-1}$  and measures  $\sigma_{t\bar{t}} = 836 \pm 27(\text{stat.}) \pm 88(\text{syst.}) \pm 100(\text{lumi.})$  pb [45].

These experimental results should be compared to the theoretical calculations at NNLO+NNLL that yield  $7.16^{+0.20}_{-0.23}$  pb for top-quark mass of  $173.3 \text{ GeV}/c^2$  [1] at  $\sqrt{s} = 1.96 \text{ TeV}$ , and for top-quark mass of  $173.2 \text{ GeV}/c^2$   $\sigma_{t\bar{t}} = 173.6^{+4.5+8.9}_{-5.9-8.9}$  pb at  $\sqrt{s} = 7 \text{ TeV}$ ,  $\sigma_{t\bar{t}} = 247.7^{+6.3+11.5}_{-8.5-11.5}$  pb at  $\sqrt{s} = 8 \text{ TeV}$ , and  $\sigma_{t\bar{t}} = 816.0^{+19.4+34.4}_{-28.6-34.4}$  pb at  $\sqrt{s} = 13 \text{ TeV}$ , at the LHC [1].

In Fig. 1, one sees the importance of  $p\bar{p}$  at Tevatron energies where the valence antiquarks in the antiprotons contribute to the dominant  $q\bar{q}$  production mechanism. At LHC energies, the dominant production mode is gluon-gluon fusion and the  $pp$ - $p\bar{p}$  difference nearly disappears. The excellent agreement of these measurements with the theory calculations is a strong validation of QCD and the soft-gluon resummation techniques employed in the calculations. The measurements reach high precision and provide stringent tests of pQCD calculations at NNLO+NNLL level including their respective PDF uncertainties.

Most of these measurements assume a  $t \rightarrow Wb$  branching ratio of 100%. CDF and DØ have made direct measurements of the  $t \rightarrow Wb$  branching ratio [47]. Comparing the number of events with 0, 1 and 2 tagged  $b$  jets in the lepton+jets channel, and also in the dilepton channel, using the known  $b$ -tagging efficiency, the ratio  $R = B(t \rightarrow Wb) / \sum_{q=d,s,b} B(t \rightarrow Wq)$  can be extracted. In  $5.4 \text{ fb}^{-1}$  of data, DØ measures  $R = 0.90 \pm 0.04$ ,  $2.5\sigma$  from unity. The currently most precise measurement was made by CMS in  $19.7 \text{ fb}^{-1}$  at  $\sqrt{s} = 8 \text{ TeV}$ . They find  $R = 1.014 \pm 0.003(\text{stat.}) \pm 0.032(\text{syst.})$  and  $R > 0.955$  at 95% C.L. [48]. A significant deviation of  $R$  from unity would imply either non-SM top-quark decay (for example a flavor-changing neutral-current decay), or a fourth generation of quarks.

Thanks to the large available event samples, the Tevatron and the LHC experiments also performed differential cross-section measurements in  $t\bar{t}$  production. Such measurements are crucial, as they allow even more stringent tests of perturbative QCD as description of the production mechanism, allow the extraction or the use of PDF fits, and enhance the sensitivity to possible new physics contributions, especially now that NNLO predictions for the main differential observables in  $t\bar{t}$  prediction have become available [51]. Furthermore, such measurements reduce the uncertainty in the description of  $t\bar{t}$  production as background in Higgs physics and searches for rare processes or beyond Standard Model physics. Differential cross-sections are typically measured by a selection of candidate events, their kinematic reconstruction and subsequent unfolding of the obtained event counts in bins of kinematic distributions in order to correct for detector resolution effects, acceptance and migration effects. In some cases a bin-by-bin unfolding is used, while other analyses use a more sophisticated techniques.

Experiments at Tevatron and LHC measure the differential cross-section with respect to the  $t\bar{t}$  invariant mass,  $d\sigma/dM_{t\bar{t}}$ . The spectra are fully corrected for detector efficiency and resolution effects and are compared to several Monte Carlo

simulations as well as selected theoretical calculations. Using  $2.7 \text{ fb}^{-1}$ , CDF measured  $d\sigma/dM_{t\bar{t}}$  in the lepton+jets channel providing sensitivity to a variety of exotic particles decaying into  $t\bar{t}$  pairs [52]. In  $9.7 \text{ fb}^{-1}$  of lepton+jets data, DØ measured the differential  $t\bar{t}$  production cross-section with respect to the transverse momentum and absolute rapidity of the top quarks as well as of the invariant mass of the  $t\bar{t}$  pair [53], which are all found to be in good agreement with the SM predictions. Also ATLAS measured the differential  $t\bar{t}$  production cross-section with respect to the top-quark transverse momentum, and of the mass, transverse momentum and rapidity of the top-quark, the antitop-quark as well as the  $t\bar{t}$  system in  $4.6 \text{ fb}^{-1}$  at  $\sqrt{s} = 7 \text{ TeV}$  in the lepton+jets channel [54–56]. The results show sensitivity to these predictions and to different sets of parton distribution functions. It is found that data is softer than all predictions for higher values of the mass of the  $t\bar{t}$  system as well as in the tail of the top-quark  $p_T$  spectrum beginning at 200 GeV, particularly in the case of the **Alpgen+Herwig** generator. The  $M_{t\bar{t}}$  spectrum is not well described by NLO+NNLL calculations and there are also disagreements between the measured  $y_{t\bar{t}}$  spectrum and the **MC@NLO+Herwig** and **POWHEG+Herwig** generators, both evaluated with the CT10 PDF set. All distributions show a preference for HERAPDF1.5 when used for the NLO QCD predictions. Recently, using  $20.3 \text{ fb}^{-1}$  of 8 TeV data, ATLAS performed a dedicated differential  $t\bar{t}$  cross section measurement of highly boosted top quarks, where the hadronically decaying top quark has a transverse momentum above 300 GeV [57]. Jet substructure techniques are employed to identify top quarks, which are reconstructed with an anti- $k_t$  jet with a radius parameters  $R = 1.0$ . The predictions of next-to-leading-order and leading-order matrix element plus parton shower Monte Carlo generators are found to generally overestimate the measured cross sections. A corresponding analysis at high transverse momentum regime for the top quarks, is performed by the CMS collaboration in  $19.7 \text{ fb}^{-1}$  at  $\sqrt{s} = 8 \text{ TeV}$  [58]. The measurement is performed for events in electron/muon plus jets final states where the hadronically decaying top quark is reconstructed as a single large-radius jet and identified as a top candidate using jet substructure techniques. The integrated cross section is measured at particle-level within a fiducial region resembling the detector-level selection as well as at parton-level. At particle-level, the cross section is measured to be  $\sigma_{t\bar{t}} = 1.28 \pm 0.09(\text{stat.} + \text{syst.}) \pm 0.10(\text{pdf}) \pm 0.09(\text{scales}) \pm 0.03(\text{lumi.})$  pb for  $p_T > 400 \text{ GeV}$ . At parton-level, it translates to  $\sigma_{t\bar{t}} = 1.44 \pm 0.10(\text{stat.} + \text{syst.}) \pm 0.13(\text{pdf}) \pm 0.15(\text{scales}) \pm 0.04(\text{lumi.})$  pb, 14% lower than the SM prediction of **POWHEG+Pythia6**. In  $5.0 \text{ fb}^{-1}$  of  $\sqrt{s} = 7 \text{ TeV}$  data in the lepton+jets and the dilepton channels, CMS measured normalised differential  $t\bar{t}$  cross-sections with respect to kinematic properties of the final-state charged leptons and jets associated to  $b$ -quarks, as well as those of the top quarks and the  $t\bar{t}$  system. The data are compared with several predictions from perturbative QCD calculations and found to be consistent [59]. Recently, in  $19.7 \text{ fb}^{-1}$

at  $\sqrt{s} = 8 \text{ TeV}$ , CMS repeated those measurements in the lepton+jets and in the dilepton channels [60]. While the overall precision is improved, no significant deviations from the Standard Model are found, yet a softer spectrum for the top quark at high  $p_T$  with respect to theoretical available predictions has been observed. This behaviour has been also observed in the all-jets final state [41].

Very recently, they also performed differential cross-section measurements in  $42 \text{ pb}^{-1}$  of single-lepton data at 13 TeV with respect to kinematic properties of the top quarks and the  $t\bar{t}$  system, as well as of the jet multiplicity in the event. The results are confronted with several predictions from pQCD and found to be consistent [61].

Further cross-section measurements are performed for  $t\bar{t}$ +heavy flavour [62] and  $t\bar{t}$ +jets production as well as the differential measurement of the jet multiplicity in  $t\bar{t}$  events [63,64]. Here, **MC@NLO+Herwig** MC is found to predict too few events at higher jet multiplicities. In addition, CMS measured the cross section ratio  $\sigma_{t\bar{t}b\bar{b}}/\sigma_{t\bar{t}jj}$  using  $19.6 \text{ fb}^{-1}$  of 8 TeV data [65]. This is of high relevance for top quark production as background to searches, for example for the ongoing search for  $t\bar{t}h$  production. Very recently, ATLAS also measured the  $t\bar{t}$  production cross section along with as the branching ratios into channels with leptons and quarks using  $4.6 \text{ fb}^{-1}$  of 7 TeV data [66]. They find agreement with the standard model at the level of a few percent.

**C.1.2 Single-top production:** Single-top quark production was first observed in 2009 by DØ [67] and CDF [68,69] at the Tevatron. The production cross section at the Tevatron is roughly half that of the  $t\bar{t}$  cross section, but the final state with a single  $W$ -boson and typically two jets is less distinct than that for  $t\bar{t}$  and much more difficult to distinguish from the background of  $W$ +jets and other sources. A comprehensive review of the first observation and the techniques used to extract the signal from the backgrounds can be found in [70].

The dominant production at the Tevatron is through  $s$ -channel and  $t$ -channel  $W$ -boson exchange. Associated production with a  $W$ -boson ( $Wt$  production) has a cross section that is too small to observe at the Tevatron. The  $t$ -channel process is  $qb \rightarrow q't$ , while the  $s$ -channel process is  $q\bar{q}' \rightarrow t\bar{b}$ . The  $s$ - and  $t$ -channel productions can be separated kinematically. This is of particular interest because potential physics beyond the Standard Model, such as fourth-generation quarks, heavy  $W$  and  $Z$  bosons, flavor-changing-neutral-currents [10], or a charged Higgs boson, would affect the  $s$ - and  $t$ -channels differently. However, the separation is difficult and initial observations and measurements at the Tevatron by both experiments were of combined  $s + t$ -channel production. The two experiments combined their measurements for maximum precision with a resulting  $s + t$ -channel production cross section of  $2.76^{+0.58}_{-0.47} \text{ pb}$  [71]. The measured value assumes a top-quark mass of  $170 \text{ GeV}/c^2$ . The mass dependence of the result comes both from the acceptance dependence and from the  $t\bar{t}$  background evaluation. Also the

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shape of discriminating topological variables is sensitive to  $m_t$ . It is therefore not necessarily a simple linear dependence but amounts to only a few tenths of picobarns over the range 170 – 175 GeV/ $c^2$ . The measured value agrees well with the theoretical calculation at  $m_t = 173$  GeV/ $c^2$  of  $\sigma_{s+t} = 3.12$  pb (including both top and anti-top production) [6,8].

Using the full Run-II data set of up to 9.7 fb $^{-1}$ , CDF and DØ have measured the  $t$ -channel single-top quark production to be  $\sigma_t = 2.25^{+0.29}_{-0.31}$  pb [72]. In the same publication, they also present the simultaneously measured  $s$ - and  $t$ -channel cross sections and the  $s+t$  combined cross section measurement resulting in  $\sigma_{s+t} = 3.30^{+0.52}_{-0.40}$  pb, without assuming the SM ratio of  $\sigma_s/\sigma_t$ . The modulus of the CKM matrix element obtained from the  $s+t$ -channel measurement is  $|V_{tb}| = 1.02^{+0.06}_{-0.05}$  and its value is used to set a lower limit of  $|V_{tb}| > 0.92$  at 95% C.L. Those results are in good agreement with the theoretical value at the mass 172.5 GeV/ $c^2$  of  $\sigma_t = 2.08 \pm 0.13$  pb [6]. It should be noted that the theory citations here list cross sections for  $t$  or  $\bar{t}$  alone, whereas the experiments measure the sum. At the Tevatron, these cross sections are equal. The theory values quoted here already include this factor of two.

Using datasets of 9.7 fb $^{-1}$  each, CDF and DØ combine their analyses and report the first observation of single-top-quark production in the  $s$ -channel, yielding  $\sigma_s = 1.29^{+0.26}_{-0.24}$  pb [73]. The probability of observing a statistical fluctuation of the background of the given size is  $1.8 \times 10^{-10}$ , corresponding to a significance of 6.3 standard deviations.

At the LHC, the  $t$ -channel cross section is expected to be more than three times as large as  $s$ -channel and  $Wt$  production, combined. Both ATLAS and CMS have measured single top production cross sections at  $\sqrt{s} = 7$  TeV in  $pp$  collisions (assuming  $m_t = 172.5$  GeV/ $c^2$  unless noted otherwise).

Using 4.59 fb $^{-1}$  of data, ATLAS measures the  $t$ -channel single-top quark cross section in the lepton plus 2 or 3 jets channel with one  $b$ -tag by fitting the distribution of a multivariate discriminant constructed with a neural network, yielding  $\sigma_t = 46 \pm 6$  pb,  $\sigma_{\bar{t}} = 23 \pm 4$  pb with a ratio  $R_t = \sigma_t/\sigma_{\bar{t}} = 2.04 \pm 0.18$  and  $\sigma_{t+\bar{t}} = 68 \pm 8$  pb, consistent with SM expectations [74]. CMS follows two approaches in 1.6 fb $^{-1}$  of lepton plus jets events. The first approach exploits the distributions of the pseudorapidity of the recoil jet and reconstructed top-quark mass using background estimates determined from control samples in data. The second approach is based on multivariate analysis techniques that probe the compatibility of the candidate events with the signal. They find  $\sigma_t = 67.2 \pm 6.1$  pb, and  $|V_{tb}| = 1.020 \pm 0.046(\text{exp.}) \pm 0.017(\text{th.})$  [76].

At  $\sqrt{s} = 8$  TeV, both experiments repeat and refine their measurements. ATLAS uses 20.3 fb $^{-1}$  by performing a combined binned maximum likelihood fit to the neural network output distribution. The measured  $t$ -channel cross-section is  $\sigma_t = 82.6 \pm 1.2(\text{stat.}) \pm 11.4(\text{syst.}) \pm 3.1(\text{pdf}) \pm 2.3(\text{lumi.})$  pb with  $|V_{tb}| = 0.97^{+0.09}_{-0.10}$  and  $|V_{tb}| > 0.78$  at 95% C.L. [77].

CMS uses 19.7 fb $^{-1}$  in the electron or muon plus jets channel, exploiting the pseudorapidity distribution of the recoil jet. They find  $\sigma_t = 53.8 \pm 1.5(\text{stat.}) \pm 4.4(\text{syst.})$  pb and  $\sigma_{\bar{t}} = 27.6 \pm 1.3(\text{stat.}) \pm 3.7(\text{syst.})$  pb, resulting in an inclusive  $t$ -channel cross section of  $\sigma_{t+\bar{t}} = 83.6 \pm 2.3(\text{stat.}) \pm 7.4(\text{syst.})$  [78]. They measure a cross section ratio of  $R_t = \sigma_t/\sigma_{\bar{t}} = 1.95 \pm 0.10(\text{stat.}) \pm 0.19(\text{syst.})$ , in agreement with the SM. The CKM matrix element  $V_{tb}$  is extracted to be  $|V_{tb}| = 0.998 \pm 0.038(\text{exp.}) \pm 0.016(\text{th.})$ .

More recently, CMS has also provided a fiducial cross section measurement for  $t$ -channel single top at  $\sqrt{s} = 8$  TeV with 19.7 fb $^{-1}$  of data in signal events with exactly one muon or electron and two jets, one of which is associated with a  $b$ -hadron. The definition of the fiducial phase space follows closely the constraints imposed by event-selection criteria and detector acceptance. The total fiducial cross section is measured using different generators at next-to-leading order plus parton-shower accuracy. Using as reference the aMC@NLO MC predictions in the four-flavour scheme a  $\sigma_t^{\text{fid}} = 3.38 \pm 0.25(\text{exp.}) \pm 0.20(\text{th.})$  pb is obtained, in good agreement with the theory predictions.

A measurement of the  $t$ -channel single top-quark cross section is also available at 13 TeV with the CMS detector, corresponding to an integrated luminosity of 42 pb $^{-1}$ . The measured cross-section is  $\sigma_t = 274 \pm 98(\text{stat.}) \pm 52(\text{syst.}) \pm 33(\text{lumi.})$  pb [79].

The  $s$ -channel production cross section is expected to be only  $4.6 \pm 0.3$  pb for  $m_t = 173$  GeV/ $c^2$  at  $\sqrt{s} = 7$  TeV [8]. The  $Wt$  process has a theoretical cross section of  $15.6 \pm 1.2$  pb [9]. This is of interest because it probes the  $Wtb$  vertex in a different kinematic region than  $s$ - and  $t$ -channel production, and because of its similarity to the associated production of a charged-Higgs boson and a top quark. The signal is difficult to extract because of its similarity to the  $t\bar{t}$  signature. Furthermore, it is difficult to uniquely define because at NLO a subset of diagrams have the same final state as  $t\bar{t}$  and the two interfere [80]. The cross section is calculated using the *diagram removal* technique [81] to define the signal process. In the diagram removal technique the interfering diagrams are removed, at the amplitude level, from the signal definition (an alternative technique, *diagram subtraction* removes these diagrams at the cross-section level and yields similar results [81]). These techniques work provided the selection cuts are defined such that the interference effects are small, which is usually the case.

Both, ATLAS and CMS, also provide evidence for the associate  $Wt$  production at  $\sqrt{s} = 7$  TeV [82,83]. ATLAS uses 2.05 fb $^{-1}$  in the dilepton plus missing  $E_T$  plus jets channel, where a template fit to the final classifier distributions resulting from boosted decision trees as signal to background separation is performed. The result is incompatible with the background-only hypothesis at the  $3.3\sigma$  ( $3.4\sigma$  expected) level, yielding  $\sigma_{Wt} = 16.8 \pm 2.9(\text{stat.}) \pm 4.9(\text{syst.})$  pb and  $|V_{tb}| = 1.03^{+0.16}_{-0.19}$  [82]. CMS uses 4.9 fb $^{-1}$  in the dilepton plus jets channel with at least one  $b$ -tag. A multivariate analysis based on kinematic properties is utilized to separate the  $t\bar{t}$  background from the signal. The

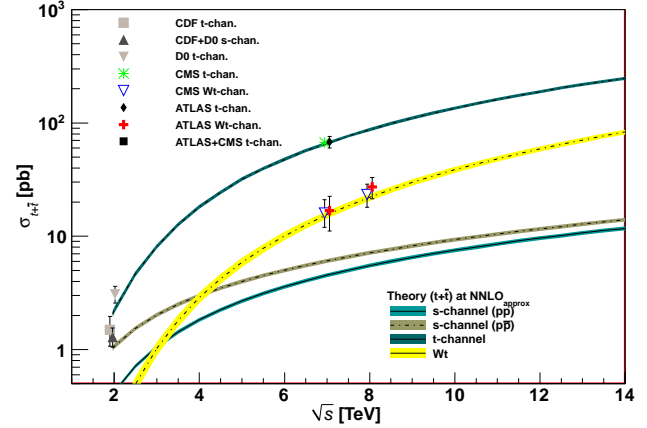
observed signal has a significance of  $4.0\sigma$  and corresponds to a cross section of  $\sigma_{Wt} = 16^{+5}_{-4}$  pb [83]. Both experiments repeated their analyses at  $\sqrt{s} = 8$  TeV. ATLAS uses  $20.3 \text{ fb}^{-1}$  to select events with one electron and one oppositely-charged muon, significant missing transverse momentum and at least one  $b$ -tagged central jet. They perform a template fit to a boosted decision tree classifier distribution and obtain  $\sigma_{Wt} = 27.2 \pm 5.8$  pb and  $|V_{tb}| = 1.10 \pm 0.12(\text{exp.}) \pm 0.03(\text{th.})$  [84], which corresponds to a  $4.2\sigma$  significance. Assuming  $|V_{tb}| \gg |V_{ts}|, |V_{td}|$  they derive  $|V_{tb}| > 0.72$  at 95% C.L. CMS uses  $12.2 \text{ fb}^{-1}$  in events with two leptons and a jet originated from a  $b$ -quark. A multivariate analysis based on kinematic properties is utilized to separate the signal and background. The  $Wt$  associate production signal is observed at the level of  $6.1\sigma$ , yielding  $\sigma_{Wt} = 23.4 \pm 5.4$  pb and  $|V_{tb}| = 1.03 \pm 0.12(\text{exp.}) \pm 0.04(\text{th.})$  [85]. They also combine their measurements and obtain  $\sigma_{Wt} = 25.0 \pm 1.4(\text{stat.}) \pm 4.4(\text{syst.}) \pm 0.7(\text{lumi.}) \text{ pb} = 25.0 \pm 4.7$  pb [86], in agreement with the NLO+NNLL expectation. They extract a 95% C.L. lower limit on the CKM matrix element of  $|V_{tb}| > 0.79$ .

At ATLAS, a search for  $s$ -channel single top quark production is performed in  $0.7 \text{ fb}^{-1}$  at 7 TeV using events containing one lepton, missing transverse energy and two  $b$ -jets. Using a cut-based analysis, an observed (expected) upper limit at 95% C.L. on the  $s$ -channel cross-section of  $\sigma_s < 26.5(20.5)$  pb is obtained [87]. In 8 TeV data, both ATLAS and CMS search for  $s$ -channel production. ATLAS uses  $20.3 \text{ fb}^{-1}$  of data with one lepton, large missing transverse momentum and exactly two  $b$ -tagged jets. They perform a maximum-likelihood fit of a discriminant based on a Matrix Element Method and optimized in order to separate single top-quark  $s$ -channel events from the main background contributions which are top-quark pair production and  $W$  boson production in association with heavy flavour jets. They find  $\sigma_s = 4.8 \pm 1.1$  pb with a signal significance of 3.2 standard deviations [88]. CMS uses  $19.3 \text{ fb}^{-1}$  and analyses leptonic decay modes by performing a likelihood fit to a multivariate discriminant as form by a Boosted Decision Tree, yielding an upper limit of  $\sigma_s < 11.5$  pb at 95% C.L. [89].

Fig. 2 provides a summary of all single top cross-section measurements at the Tevatron and the LHC as a function of the center-of-mass energy. All cross-section measurements are very well described by the theory calculation within their uncertainty.

Thanks to the large statistics now available at the LHC, both CMS and ATLAS experiments also performed differential cross-section measurements in single-top  $t$ -channel production [74], [98]. Such measurements are extremely useful as they test our understanding of both QCD and EW top-quark interactions.

The CMS collaboration has measured differential single top quark  $t$ -channel production cross sections as functions of the transverse momentum and the absolute value of the rapidity of the top quark. The analysis is performed in the leptonic decay channels of the top quark, with either a muon or an electron in the final state, using data collected with the CMS



**Figure 2:** Measured and predicted single top production cross sections from Tevatron energies in  $p\bar{p}$  collisions to LHC energies in  $pp$  collisions. Tevatron data points at  $\sqrt{s} = 1.96$  TeV are from Refs. [90,91] and [92]. The ATLAS and CMS data points at  $\sqrt{s} = 7$  TeV are from Refs. [75,82,87,93] and [76,83,94], respectively. The ones at  $\sqrt{s} = 8$  TeV are from Refs. [84,96] and [95,96,97]. Theory curves are generated using [6,8,9].

experiment at the LHC at  $\sqrt{s} = 8$  TeV and corresponding to an integrated luminosity of  $19.7 \text{ fb}^{-1}$ . Artificial neural networks are used to discriminate the signal process from the various background contributions. The results are found to agree with predictions from Monte Carlo generators [98]. Using the same data set and under the assumption that the spin analyzing power of a charged lepton is 100% as predicted in the SM, they are also able to measure the polarization of the top quark  $P_t = 0.82 \pm 0.12(\text{stat.}) \pm 0.32(\text{syst.})$  [99].

**C.1.3 Top-Quark Forward-Backward & Charge Asymmetry:** A forward-backward asymmetry in  $t\bar{t}$  production arises starting at order  $\alpha_S^3$  in QCD from the interference between the Born amplitude  $q\bar{q} \rightarrow t\bar{t}$  with 1-loop box production diagrams and between diagrams with initial- and final-state gluon radiation. The asymmetry,  $A_{FB}$ , is defined by

$$A_{FB} = \frac{N(\Delta y > 0) - N(\Delta y < 0)}{N(\Delta y > 0) + N(\Delta y < 0)} \quad (2)$$

where  $\Delta y = y_t - y_{\bar{t}}$  is the rapidity difference between the top- and the anti-top quark. Calculations at  $\alpha_S^3$  predict a small  $A_{FB}$  at the Tevatron. The most recent calculations up to order  $\alpha_S^4$ , including electromagnetic and electroweak corrections, yield a predicted asymmetry of  $(\approx 9.5 \pm 0.7)\%$  [100]. This is about 10% higher than the previous calculation at NLO [101,102], and improves the agreement with experiment.

Both, CDF and DØ, measured asymmetry values in excess of the SM prediction, fueling speculation about exotic production mechanisms (see, for example, [103] and references therein). The first measurement of this asymmetry by DØ in  $0.9 \text{ fb}^{-1}$  [104] found an asymmetry at the detector level of  $(12 \pm 8)\%$ . The first CDF measurement in  $1.9 \text{ fb}^{-1}$  [105] yielded  $(24 \pm 14)\%$  at parton level. Both values were higher, though

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statistically consistent with the SM expectation. With the addition of more data, the uncertainties have been reduced, and the central values, if somewhat smaller, have remained consistent with the first measurements. At the same time, the improved calculations from theory have increased the predicted asymmetry values to the point where the discrepancy is no longer statistically significant. The most recent measurement from  $D\bar{O}$  using the full Tevatron dataset of  $9.7 \text{ fb}^{-1}$  finds an asymmetry in lepton+jet events, corrected for detector acceptance and resolution, of  $(10.6 \pm 3.0)\%$  [106] in good agreement with the prediction. Using the same dataset, the  $D\bar{O}$  measurement in dilepton events and assuming SM top polarization is  $17.5 \pm 5.6(\text{stat.}) \pm 3.1(\text{syst.})\%$  [107]. Combining the lepton+jets with the dilepton gives  $11.8 \pm 2.5(\text{stat.}) \pm 1.3(\text{syst.})\%$ .

From CDF, the most recent measurement in lepton+jets uses  $9.4 \text{ fb}^{-1}$ , and finds  $(16.4 \pm 4.7)\%$  [108]. This measurement has now been combined with an asymmetry measured in dilepton events using  $9.1 \text{ fb}^{-1}$  [109]. The asymmetry reported for dilepton events is  $(12 \pm 13)\%$ , and the combined asymmetry is  $(16.0 \pm 4.5)\%$ , which is about  $1.5\sigma$  above the NNLO prediction.

Both experiments have measured  $A_{FB}$  as a function of  $M_{t\bar{t}}$ , the  $t\bar{t}$  invariant mass and in bins of  $|\Delta y|$  [108,106]. The experiments see, and theory predicts, a positive slope in  $A_{FB}$  with increasing  $M_{t\bar{t}}$  and  $|\Delta y|$ . The slopes seen in the CDF data remain larger than the theoretical expectation, while the  $D\bar{O}$  data are in good agreement with the latest theoretical calculation [100].

At the LHC, where the dominant  $t\bar{t}$  production mechanism is the charge-symmetric gluon-gluon fusion, the measurement is more difficult. For the sub-dominant  $q\bar{q}$  production mechanism, the symmetric  $pp$  collision does not define a forward and backward direction. Instead, the charge asymmetry,  $A_C$ , is defined in terms of a positive versus a negative  $t - \bar{t}$  rapidity difference

$$A_C = \frac{N(\Delta y > 0) - N(\Delta y < 0)}{N(\Delta y > 0) + N(\Delta y < 0)} \quad (3)$$

Both CMS and ATLAS have measured  $A_C$  in the LHC dataset. Using lepton+jets events in  $4.7 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$ , ATLAS measures  $A_C = (0.6 \pm 1.0)\%$  [110]. More recently, ATLAS has reported on the same measurement performed at  $\sqrt{s} = 8 \text{ TeV}$  with at  $20.3 \text{ fb}^{-1}$  of data. The result is  $A_C = (0.009 \pm 0.005)$  [111]. CMS, in  $5.0(19.7) \text{ fb}^{-1}$  of  $\sqrt{s} = 7(8) \text{ TeV}$  data uses lepton+jets events to measure  $A_C = (0.4 \pm 1.5)\%$  ( $A_C = (0.33 \pm 0.26(\text{stat.}) \pm 0.33(\text{syst.}))\%$ ) [112,113]. Both measurements are consistent with the SM expectations of  $A_C = 1.23 \pm 0.05\%$  at  $\sqrt{s} = 7 \text{ TeV}$  and  $1.11 \pm 0.04\%$  at  $\sqrt{s} = 8 \text{ TeV}$  [102], although the uncertainties are still too large for a precision test. In their 7 and 8 TeV analyses ATLAS and CMS also provide differential measurements as a function of  $M_{t\bar{t}}$  and the transverse momentum  $p_T$  and rapidity  $y$  of the  $t\bar{t}$  system. In a recent work [114] the CMS collaboration has provided the result of  $A_C = -0.0035 \pm 0.0072(\text{stat.}) \pm 0.0031(\text{syst.})$  obtained in the fiducial phase space of top quark pair production.

Another avenue for measuring the forward-backward and charge asymmetries that has recently been exploited by the experiments is given by the measurement of the pseudorapidity distributions of the charged leptons resulting from  $t\bar{t}$  decay. Although the expected asymmetry is smaller, this technique does not require the reconstruction of the top-quark direction. Single-lepton asymmetries,  $A_{FB}^\ell$ , are defined by  $q \times \eta$ , and dilepton asymmetries,  $A^{\ell\ell}$ , by the sign of  $\Delta\eta$ , where  $q$  and  $\eta$  are the charge and pseudorapidity of the lepton and  $\Delta\eta = \eta_{\ell^+} - \eta_{\ell^-}$ .  $D\bar{O}$  has measured  $A_{FB}^\ell$  in  $9.7 \text{ fb}^{-1}$  of lepton+jets events, and finds a value of  $(4.2 \pm 2.3^{+1.7}_{-2.0})\%$  [115], consistent with an expectation of  $(3.8 \pm 0.6)\%$  [102]. A measurement by  $D\bar{O}$  using dilepton events in the same dataset [116] yields  $A^{\ell\ell} = (12.3 \pm 5.4 \pm 1.5)\%$ , compared to the expectation of  $(4.8 \pm 0.4)\%$  [102], and  $A_{FB}^\ell = 4.4 \pm 3.7 \pm 1.1$ . The combination of the results for  $A_{FB}^\ell$  in the single lepton and dilepton channels by  $D\bar{O}$  yields  $(4.2 \pm 2.0 \pm 1.4)\%$ . CDF, in  $9.4 \text{ fb}^{-1}$  of Tevatron data measures [117]  $A_{FB}^\ell = (9.4^{+3.2}_{-2.9})\%$ . As in the  $D\bar{O}$  case, this is larger than the SM expectation, but less than two standard deviations away.

At the LHC, both ATLAS and CMS have now measured leptonic asymmetries. ATLAS, in  $4.6 \text{ fb}^{-1}$  of  $\sqrt{s} = 7 \text{ TeV}$  data, has measured  $A^{\ell\ell} = (2.4 \pm 1.5 \pm 0.9)\%$  in dilepton events [118]. Using a neutrino weighting technique in the same dataset to reconstruct the top quarks, ATLAS measures  $A_C = (2.1 \pm 2.5 \pm 1.7)\%$ . CMS, in  $5.0 \text{ fb}^{-1}$  of  $\sqrt{s} = 7 \text{ TeV}$  data, uses dilepton events to measure  $A_C = (1.0 \pm 1.5 \pm 0.6)\%$ , where a matrix weighting technique is used to reconstruct the top quarks, and  $A^{\ell\ell} = (0.9 \pm 1.0 \pm 0.6)\%$  [119]. An earlier result using lepton+jets events from the same CMS dataset found  $A_C = (0.4 \pm 1.0 \pm 1.1)\%$  [112]. These results are all consistent, within their large uncertainties, with the SM expectations of  $A^{\ell\ell} = (0.70 \pm 0.03)\%$  and  $A_C = (1.23 \pm 0.05)\%$  [102].

A model-independent comparison of the Tevatron and LHC results is made difficult by the differing  $t\bar{t}$  production mechanisms at work at the two accelerators and by the symmetric nature of the  $pp$  collisions at the LHC. Given a particular model of BSM physics, a comparison can be obtained through the resulting asymmetry predicted by the model at the two machines, see for example [120].

## C.2 Top-Quark Properties

**C.2.1 Top-Quark Mass Measurements:** The most precisely studied property of the top quark is its mass. The top-quark mass has been measured in the lepton+jets, the dilepton, and the all-jets channel by all four Tevatron and LHC experiments. The latest and/or most precise results are summarized in Table 1. The lepton+jets channel yields the most precise single measurements because of good signal to background ratio (in particular after  $b$ -tagging) and the presence of only a single neutrino in the final state. The momentum of a single neutrino can be reconstructed (up to a quadratic ambiguity) via the missing  $E_T$  measurement and the constraint that the

lepton and neutrino momenta reconstruct to the known  $W$  boson mass. In the large data samples available at the LHC, measurements in the dilepton channel can be competitive and certainly complementary to those in the lepton+jets final state.

A large number of techniques have now been applied to measuring the top-quark mass. The original ‘template method’ [121], in which Monte Carlo templates of reconstructed mass distributions are fit to data, has evolved into a precision tool in the lepton+jets channel, where the systematic uncertainty due to the jet energy scale (JES) uncertainty is controlled by a simultaneous, *in situ* fit to the  $W \rightarrow jj$  hypothesis [122]. All the latest measurements in the lepton+jets and the all-jets channels use this technique in one way or another. In  $4.6 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$  in the lepton+jets channel, ATLAS achieves a total uncertainty of 0.73% with a statistical component of 0.44% [123]. The measurement is based on a 3-dimensional template fit, determining the top-quark mass, the global jet energy scale and a  $b$ -to-light jet energy scale factor. In  $19.7 \text{ fb}^{-1}$  of  $\sqrt{s} = 8 \text{ TeV}$  data, CMS achieves a total uncertainty of 0.45% with a statistical component of 0.11% [124].

The template method is complemented by the ‘matrix element’ method. This method was first applied by the DØ Collaboration [125], and is similar to a technique originally suggested by Kondo *et al.* [126] and Dalitz and Goldstein [127]. In the matrix element method a probability for each event is calculated as a function of the top-quark mass, using a LO matrix element for the production and decay of  $t\bar{t}$  pairs. The *in situ* calibration of dijet pairs to the  $W \rightarrow jj$  hypothesis is now also used with the matrix element technique to constrain the jet energy scale uncertainty. The latest measurement with this technique from DØ in the lepton+jets channel uses the full Tevatron dataset of  $9.7 \text{ fb}^{-1}$  and yields an uncertainty of about 0.43% [128].

In the dilepton channel, the signal to background is typically very good, but reconstruction of the mass is non-trivial because there are two neutrinos in the final state, yielding a kinematically unconstrained system. A variety of techniques have been developed to handle this. An analytic solution to the problem has been proposed [129], but this has not yet been used in the mass measurement. One of the most precise measurements in the dilepton channel comes from using the invariant mass of the charged lepton and  $b$ -quark system ( $M_{\ell b}$ ), which is sensitive to the top-quark mass and avoids the kinematic difficulties of the two-neutrino final state. In  $4.6 \text{ fb}^{-1}$  of  $\sqrt{s} = 7 \text{ TeV}$  data, ATLAS has measured the top-quark mass in the dilepton channel to a precision of 0.81% using a template fit to the  $M_{\ell b}$  distribution [123]. A similar measurement has been also provided by CMS [130], giving a precision of 0.75%. The other dilepton-channel measurement of similar precision comes from  $19.7 \text{ fb}^{-1}$  of CMS data at  $\sqrt{s} = 8 \text{ TeV}$  [131] using a so-called analytical matrix weighting technique (AMWT) in which each event is fit many times to a range of top-quark masses and each fit is assigned a weight, from the PDFs, given by the inferred kinematics of the initial state partons, and from

the probability of the observed charged lepton energies for the top-quark mass in question.

Several other techniques can also yield precise measurements in the dilepton channel. In the neutrino weighting technique, similar to AMWT above, a weight is assigned by assuming a top-quark mass value and applying energy-momentum conservation to the top-quark decay, resulting in up to four possible pairs of solutions for the neutrino and anti-neutrino momenta. The missing  $E_T$  calculated in this way is then compared to the observed missing  $E_T$  to assign a weight [132]. A recent CDF result, using the full  $9.1 \text{ fb}^{-1}$  dataset achieves a precision of 1.8% using a combination of neutrino weighting and an “alternative mass”, which is insensitive to the jet energy scale [133]. The alternative mass depends on the angles between the leptons and the leading jets and the lepton four-momenta.

In the all-jets channel there is no ambiguity due to neutrino momenta, but the signal to background is significantly poorer due to the severe QCD multijets background. The emphasis therefore has been on background modeling, and reduction through event selection. The most recent measurement in the all-jets channel, by CMS in  $18.2 \text{ fb}^{-1}$  of  $\sqrt{s} = 8 \text{ TeV}$  data [134], uses an ideogram and a 2-dimensional simultaneous fit for  $m_t$  and the jet energy scale to extract the top-quark mass and achieves a precision of 0.53%. A recent measurement from ATLAS [135] uses the template method in the all-hadronic channel, also with an *in situ* fit to the  $W \rightarrow jj$  hypothesis, yielding a measurement with 1.9% precision in  $4.6 \text{ fb}^{-1}$  of data. A measurement from CDF in  $9.3 \text{ fb}^{-1}$  uses similar two-dimensional template fit and achieves a precision of 1.1% [136].

A dominant systematic uncertainty in these methods is the understanding of the jet energy scale, and so several techniques have been developed that have little sensitivity to the jet energy scale uncertainty. In addition to Reference [133] mentioned above, these include the measurement of the top-quark mass using the following techniques: Fitting of the lepton  $p_T$  spectrum of candidate events [137]; fitting of the transverse decay length of the  $b$ -jet ( $L_{xy}$ ) [138]; fitting the invariant mass of a lepton from the  $W$ -decay and a muon from the semileptonic  $b$  decay [139].

Several measurements have now been made in which the top-quark mass is extracted from the measured cross section using the theoretical relationship between the mass and the production cross section. These determinations make use of predictions calculated at higher orders, where the top mass enters as an input parameter defined in a given scheme. At variance with the usual methods, which involve the kinematic properties of the final states and therefore the pole mass, this approach allows to directly determine a short-distance mass, such as the  $\overline{\text{MS}}$  mass [140]. With an alternative method ATLAS recently extracted the top-quark pole mass using  $t\bar{t}$  events with at least one additional jet, basing the measurement on the relationship between the differential rate of gluon radiation and the mass of the quark [141].



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Each of the experiments has produced a measurement combining its various results. The combined measurement from CMS with up to  $19.7 \text{ fb}^{-1}$  of data achieves statistical and systematic uncertainties of 0.06% and 0.38%, respectively [142]. The combined measurement from ATLAS, with  $4.6 \text{ fb}^{-1}$  yields statistical and systematic uncertainties of 0.28% and 0.45%, respectively [123]. CDF has combined measurements with up to  $9.3 \text{ fb}^{-1}$  [143] and achieves a statistical precision of 0.33% and a systematic uncertainty of 0.43%. DØ achieves a 0.33% statistical+JES and a 0.28% systematic uncertainty by combining results in  $9.7 \text{ fb}^{-1}$  [144].

Combined measurements from the Tevatron experiments and from the LHC experiments take into account the correlations between different measurements from a single experiment and between measurements from different experiments. The Tevatron average [145], using up to  $9.7 \text{ fb}^{-1}$  of data, now has a precision of 0.37%. The LHC combination, using up to  $4.9 \text{ fb}^{-1}$  of data, has a precision of 0.56% [146], where more work on systematic uncertainties is required. The first Tevatron-LHC combination has now been released, combining the results of all four experiments, using the full Tevatron dataset and the  $\sqrt{s} = 7 \text{ TeV}$  LHC data, with a resulting precision of 0.44% [3]

The direct measurements of the top-quark mass, such as those shown in Table 1, strictly speaking, is the corresponding parameter used in the Monte Carlo generators. The relation between the parameter in the Monte Carlo generator and the pole mass is affected by non-perturbative contributions, which could be order  $1 \text{ GeV}/c^2$  [147], i.e., comparable to the measurement uncertainty.

With the discovery of a Higgs boson at the LHC with a mass of about  $126 \text{ GeV}/c^2$  [148,149], the precision measurement of the top-quark mass takes a central role in the question of the stability of the electroweak vacuum because top-quark radiative corrections tend to drive the Higgs quartic coupling,  $\lambda$ , negative, potentially leading to an unstable vacuum. A recent calculation at NNLO [150] leads to the conclusion of vacuum stability for a Higgs mass satisfying  $M_H \geq 129.4 \pm 5.6 \text{ GeV}/c^2$  [151]. Given the uncertainty, a Higgs mass of  $126 \text{ GeV}/c^2$  satisfies the limit, but the central values of the Higgs and top-quark masses put the electroweak vacuum squarely in the metastable region. The uncertainty is dominated by the precision of the top-quark mass measurement and its interpretation as the pole mass. For more details, see the Higgs boson review in this volume.

As a test of the CPT-symmetry, the mass difference of top- and antitop-quarks  $\Delta m_t = m_t - m_{\bar{t}}$ , which is expected to be zero, can be measured. CDF measures the mass difference in  $8.7 \text{ fb}^{-1}$  of 1.96 TeV data in the lepton+jets channel using a template method to find  $\Delta m_t = -1.95 \pm 1.11(\text{stat.}) \pm 0.59(\text{syst.}) \text{ GeV}/c^2$  [152] while DØ uses  $3.6 \text{ fb}^{-1}$  of lepton+jets events and the matrix element method with at least one *b*-tag. They find  $\Delta m_t = 0.8 \pm 1.8(\text{stat.}) \pm 0.5(\text{syst.}) \text{ GeV}/c^2$  [153]. In  $4.7 \text{ fb}^{-1}$  of 7 TeV data, ATLAS measures the mass difference in lepton+jets events with a double *b*-tag requirement and hence very low background to find  $\Delta m_t = 0.67 \pm 0.61(\text{stat.}) \pm$

**Table 1:** Measurements of top-quark mass from Tevatron and LHC.  $\int \mathcal{L} dt$  is given in  $\text{fb}^{-1}$ . The results shown are mostly preliminary (not yet submitted for publication as of August 2015); for a complete set of published results see the Listings. Statistical uncertainties are listed first, followed by systematic uncertainties.

$m_t \text{ (GeV}/c^2\text{)}$	Source	$\int \mathcal{L} dt$	Ref.	Channel
$172.99 \pm 0.48 \pm 0.78$	ATLAS	4.6	[123]	$\ell$ +jets+ $\ell\ell$
$172.04 \pm 0.19 \pm 0.75$	CMS	19.7	[124]	$\ell$ +jets
$172.47 \pm 0.17 \pm 1.40$	CMS	19.7	[131]	$\ell\ell$
$172.32 \pm 0.25 \pm 0.59$	CMS	19.7	[134]	All jets
$174.34 \pm 0.37 \pm 0.52$	CDF,DØ (I+II)	$\leq 9.7$	[145]	publ. or prelim.
$173.34 \pm 0.27 \pm 0.71$	Tevatron+LHC	$\leq 8.7 + \leq 4.9$	[3]	publ. or prelim.

$0.41(\text{syst.}) \text{ GeV}/c^2$  [154]. CMS measures the top-quark mass difference in  $5 \text{ fb}^{-1}$  of 7 TeV data in the lepton+jets channel and finds  $\Delta m_t = -0.44 \pm 0.46(\text{stat.}) \pm 0.27(\text{syst.}) \text{ GeV}/c^2$  [155]. They repeat this measurement with  $18.9 \text{ fb}^{-1}$  of 8 TeV data to find  $\Delta m_t = -0.27 \pm 0.20(\text{stat.}) \pm 0.12(\text{syst.}) \text{ GeV}/c^2$  [156]. All measurements are consistent with the SM expectation.

## C.2.2 Top-Quark Spin Correlations, Polarization, and Width:

One of the unique features of the top quark is that it decays before its spin can be flipped by the strong interaction. Thus the top-quark polarization is directly observable via the angular distribution of its decay products. Hence, it is possible to define and measure observables sensitive to the top-quark spin and its production mechanism. Although the top- and antitop-quarks produced by strong interactions in hadron collisions are essentially unpolarized, the spins of *t* and  $\bar{t}$  are correlated. For QCD production at threshold, the  $t\bar{t}$  system is produced in a  $^3S_1$  state with parallel spins for  $q\bar{q}$  annihilation or in a  $^1S_0$  state with antiparallel spins for gluon-gluon fusion. Hence, the situations at the Tevatron and at the LHC are somewhat complementary. However, at the LHC production of  $t\bar{t}$  pairs at large invariant mass occurs primarily via fusion of gluons with opposite helicities, and the  $t\bar{t}$  pairs so produced have parallel spins as in production at the Tevatron via  $q\bar{q}$  annihilation. The direction of the top-quark spin is 100% correlated to the angular distributions of the down-type fermion (charged leptons or *d*-type quarks) in the decay. The joint angular distribution [157–159]

$$\frac{1}{\sigma} \frac{d^2\sigma}{d(\cos\theta_+)d(\cos\theta_-)} = \frac{1 + \kappa \cdot \cos\theta_+ \cdot \cos\theta_-}{4}, \quad (4)$$

where  $\theta_+$  and  $\theta_-$  are the angles of the daughters in the top-quark rest frame with respect to a particular spin quantization axis, is a very sensitive observable. The maximum value for  $\kappa$ , 0.782 at NLO at the Tevatron [160], is found in the off-diagonal basis [157], while at the LHC the value at NLO is 0.326 in the helicity basis [160]. In place of  $\kappa$ ,  $A\alpha_+\alpha_-$  is often used, where



$\alpha_i$  is the spin analyzing power, and  $A$  is the spin correlation coefficient, defined as

$$A = \frac{N(\uparrow\uparrow) + N(\downarrow\downarrow) - N(\uparrow\downarrow) - N(\downarrow\uparrow)}{N(\uparrow\uparrow) + N(\downarrow\downarrow) + N(\uparrow\downarrow) + N(\downarrow\uparrow)}, \quad (5)$$

where the first arrow represents the direction of the top-quark spin along a chosen quantization axis, and the second arrow represents the same for the antitop-quark. The spin analyzing power  $\alpha_i$  is +0.998 for positively charged leptons, -0.966 for down-type quarks from  $W$  decays, and -0.393 for bottom quarks [161]. The sign of  $\alpha$  flips for the respective antiparticles. The spin correlation could be modified by a new  $t\bar{t}$  production mechanism such as through a  $Z'$  boson, Kaluza-Klein gluons, or a Higgs boson.

CDF used  $5.1 \text{ fb}^{-1}$  in the dilepton channel to measure the correlation coefficient in the beam axis [162]. The measurement was made using the expected distributions of  $(\cos\theta_+, \cos\theta_-)$  and  $(\cos\theta_b, \cos\theta_{\bar{b}})$  of the charged leptons or the  $b$ -quarks in the  $t\bar{t}$  signal and background templates to calculate a likelihood of observed reconstructed distributions as a function of assumed  $\kappa$ . They determined the 68% confidence interval for the correlation coefficient  $\kappa$  as  $-0.52 < \kappa < 0.61$  or  $\kappa = 0.04 \pm 0.56$  assuming  $m_t = 172.5 \text{ GeV}/c^2$ .

CDF also analyzed lepton+jets events in  $5.3 \text{ fb}^{-1}$  [163] assuming  $m_t = 172.5 \text{ GeV}/c^2$ . They form three separate templates - the same-spin template, the opposite-spin template, and the background template for the 2-dimensional distributions in  $\cos(\theta_l)\cos(\theta_d)$  vs.  $\cos(\theta_l)\cos(\theta_b)$ . The fit to the data in the helicity basis returns an opposite helicity fraction of  $F_{OH} = 0.74 \pm 0.24(\text{stat.}) \pm 0.11(\text{syst.})$ . Converting this to the spin correlation coefficient yields  $\kappa_{\text{helicity}} = 0.48 \pm 0.48(\text{stat.}) \pm 0.22(\text{syst.})$ . In the beamline basis, they find an opposite spin fraction of  $F_{OS} = 0.86 \pm 0.32(\text{stat.}) \pm 0.13(\text{syst.})$  which can be converted into a correlation coefficient of  $\kappa_{\text{beam}} = 0.72 \pm 0.64(\text{stat.}) \pm 0.26(\text{syst.})$ .

DØ performed a measurement of the ratio  $f$  of events with correlated  $t$  and  $\bar{t}$  spins to the total number of  $t\bar{t}$  events in  $5.3 \text{ fb}^{-1}$  in the lepton+jets channel using a matrix element technique [164]. The SM expectation is  $f = 1$ . From 729 events, they obtain  $f_{\text{exp.}} = 1.15^{+0.42}_{-0.43}(\text{stat.} + \text{syst.})$  and can exclude values of  $f < 0.420$  at the 95% C.L. In the dilepton channel [165], they also use a matrix element method and can exclude at the 97.7% C.L. the hypothesis that the spins of the  $t$  and  $\bar{t}$  are uncorrelated. The combination [164] yields  $f_{\text{exp.}} = 0.85 \pm 0.29(\text{stat.} + \text{syst.})$  and a  $t\bar{t}$  production cross section which is in good agreement with the SM prediction and previous measurements. For an expected fraction of  $f = 1$ , they can exclude  $f < 0.481$  at the 95% C.L. For the observed value of  $f_{\text{exp.}} = 0.85$ , they can exclude  $f < 0.344(0.052)$  at the 95(99.7)% C.L. The observed fraction  $f_{\text{exp.}}$  translates to a measured asymmetry value of  $A_{\text{exp.}} = 0.66 \pm 0.23(\text{stat.} + \text{syst.})$ . They therefore obtained the first evidence of SM spin correlation at 3.1 standard deviations.

Using  $5.4 \text{ fb}^{-1}$  of data, DØ measures the correlation in the dilepton channel also from the angles of the two leptons in the  $t$  and  $\bar{t}$  rest frames, yielding a correlation strength  $C = 0.10 \pm 0.45$  [166] ( $C$  is equivalent to negative  $\kappa$  in Eq. 4), in agreement with the NLO QCD prediction, but also in agreement with the no correlation hypothesis.

Spin correlations have now been conclusively measured at the LHC by both the ATLAS and CMS collaborations. In the dominant gluon fusion production mode for  $t\bar{t}$  pairs at the LHC, the angular distribution between the two leptons in  $t\bar{t}$  decays to dileptons is sensitive to the degree of spin correlation [167].

The ATLAS collaboration has measured spin correlations in  $t\bar{t}$  production at  $\sqrt{s} = 7 \text{ TeV}$  using  $4.6 \text{ fb}^{-1}$  of data. Candidate events are selected in the dilepton and lepton plus jets topologies. Four observables are used to extract the spin correlation: The difference,  $\Delta\phi$  in azimuthal angle between the two charged leptons in dilepton events or the lepton and down-quark or bottom-quark candidate from the hadronic  $W$ -decay; An observable based on the ratio matrix elements with and without spin correlation; The double differential distribution of Eq. 4 in two different bases. The most sensitive measurement comes from using  $\Delta\phi$  in dilepton events and results in  $f_{\text{SM}} = 1.19 \pm 0.09 \pm 0.18$ . Using the helicity basis as the quantization axis, the strength of the spin correlation between the top- and antitop-quark is measured to be  $A_{\text{helicity}}^{\text{exp.}} = 0.37 \pm 0.03 \pm 0.06$  [168], which is in agreement with the NLO prediction of about 0.31 [169]. Using the same events but converting  $f_{\text{exp.}}$  into  $A_{\text{maximal}}^{\text{exp.}}$  yields  $A_{\text{maximal}}^{\text{exp.}} = 0.52 \pm 0.04 \pm 0.08$ , to be compared to the NLO prediction of 0.44. In a similar analysis using  $20.3 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 8 \text{ TeV}$ , ATLAS measures  $f_{\text{SM}} = 1.20 \pm 0.05(\text{stat.}) \pm 0.13(\text{syst.})$ , corresponding to  $A_{\text{helicity}}^{\text{exp.}} = 0.38 \pm 0.04$  [170], which compares well to the SM expectation of  $A_{\text{helicity}}^{\text{SM}} = 0.318 \pm 0.005$  [169].

The CMS collaboration uses angular asymmetry variables in dilepton events, unfolded to the parton level. The most sensitive measurement is made using

$$A_{\Delta\phi} = \frac{N(\Delta\phi_{\ell+\ell-} > \pi/2) - N(\Delta\phi_{\ell+\ell-} < \pi/2)}{N(\Delta\phi_{\ell+\ell-} > \pi/2) + N(\Delta\phi_{\ell+\ell-} < \pi/2)}, \quad (6)$$

In  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$ , CMS measures  $A_{\Delta\phi} = 0.113 \pm 0.010 \pm 0.006 \pm 0.012$  [171], where the uncertainties are statistical, systematic, and due to the reweighting of the top  $p_T$  in the Monte Carlo to match data. A recent CMS result in  $\mu$  plus jets events in  $19.7 \text{ fb}^{-1}$  of  $\sqrt{s} = 8 \text{ TeV}$  data uses a matrix-element technique to extract  $f_{\text{exp.}} = 0.72 \pm 0.09^{+0.15}_{-0.13}$ , corresponding to  $A_{\text{helicity}}^{\text{exp.}} = 0.22 \pm 0.03^{+0.05}_{-0.04}$  [172]. Corresponding results obtained by studying the dilepton final state, also show consistency with the SM expectations [173].

Measurements of the polarization of top quarks in  $t\bar{t}$  production at  $\sqrt{s} = 7 \text{ TeV}$  have been made by both ATLAS and CMS. In  $4.7 \text{ fb}^{-1}$  of data, ATLAS measures the product of the leptonic spin-analyzing power ( $\alpha_\ell$ ) and the top quark polarization. The measurement is made in one or two lepton final states, assuming that the polarization is introduced

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by a CP-conserving (CPC) or maximally CP-violating (CPV) process. The results are  $\alpha_\ell P_{CPC} = -0.035 \pm 0.014 \pm 0.037$  and  $\alpha_\ell P_{CPV} = 0.020 \pm 0.016^{+0.013}_{-0.017}$  [174], where the uncertainties are statistical and systematic, respectively. The CMS measurement is made with  $5.0 \text{ fb}^{-1}$  of dilepton events. The polarization is extracted through an asymmetry,  $A_P$ , in the angular distribution of the two leptons,  $A_P$ , defined as

$$A_P = \frac{N(\cos \theta_\ell^* > 0) - N(\cos \theta_\ell^* < 0)}{N(\cos \theta_\ell^* > 0) + N(\cos \theta_\ell^* < 0)}, \quad (7)$$

where  $\theta^*$  is the angle of the charged lepton in the rest frame of its parent top quark or antiquark. The polarization,  $P$  in the helicity basis is given by  $P = 2A_P$ . After unfolding to the parton level, the measurement yields  $A_P = 0.005 \pm 0.013 \pm 0.014 \pm 0.008$  [171], where the uncertainties are, respectively, statistical, systematic, and from top-quark  $p_T$  reweighting. Both the ATLAS and CMS results are consistent with the SM expectation of negligible polarization.

Observation of top-quark spin correlations requires a top-quark lifetime less than the spin decorrelation timescale [175]. The top-quark width, inversely proportional to its lifetime, is expected to be of order  $1 \text{ GeV}/c^2$  (Eq. 1). The sensitivity of current experiments does not approach this level in direct measurements. Nevertheless, several measurements have been made.

CDF presents a direct measurement of the top-quark width in the lepton+jets decay channel of  $t\bar{t}$  events from a data sample corresponding to  $8.7 \text{ fb}^{-1}$  of integrated luminosity. The top-quark mass and the mass of the hadronically decaying  $W$  boson that comes from the top-quark decay are reconstructed for each event and compared with templates of different top-quark widths ( $\Gamma_t$ ) and deviations from nominal jet energy scale ( $\Delta JES$ ) to perform a simultaneous fit for both parameters, where  $\Delta JES$  is used for the *in situ* calibration of the jet energy scale. By applying a Feldman-Cousins approach, they establish an upper limit at 95% C.L. of  $\Gamma_t < 6.38 \text{ GeV}$  and a two-sided 68% C.L. interval of  $1.10 \text{ GeV} < \Gamma_t < 4.05 \text{ GeV}$ , corresponding to a lifetime interval of  $1.6 \times 10^{-15} < \tau_{top} < 6.0 \times 10^{-25}$  [176], consistent with the SM prediction. For comparison, a typical hadronization timescale is an order of magnitude larger than these limits.

The total width of the top-quark can also be determined from the partial decay width  $\Gamma(t \rightarrow Wb)$  and the branching fraction  $B(t \rightarrow Wb)$ . DØ obtains  $\Gamma(t \rightarrow Wb)$  from the measured  $t$ -channel cross section for single top-quark production in  $5.4 \text{ fb}^{-1}$ , and  $B(t \rightarrow Wb)$  is extracted from a measurement of the ratio  $R = B(t \rightarrow Wb)/B(t \rightarrow Wq)$  in  $t\bar{t}$  events in lepton+jets channels with 0, 1 and 2 b-tags. Assuming  $B(t \rightarrow Wq) = 1$ , where  $q$  includes any kinematically accessible quark, the result is:  $\Gamma_t = 2.00^{+0.47}_{-0.43} \text{ GeV}$  which translates to a top-quark lifetime of  $\tau_t = (3.29^{+0.90}_{-0.63}) \times 10^{-25} \text{ s}$ . Assuming a high mass fourth generation  $b'$  quark and unitarity of the four-generation quark-mixing matrix, they set the first upper limit on  $|V_{tb'}| < 0.59$  at 95% C.L. [177]. A similar analysis has performed by CMS in

$19.7 \text{ fb}^{-1}$  of  $\sqrt{s} = 8 \text{ TeV}$  data. It provides a better determination of the total width with respect to the measurement by DØ giving  $\Gamma_t = 1.36 \pm 0.02(\text{stat.})^{+0.14}_{-0.11}(\text{syst.}) \text{ GeV}$  [178].

**C.2.3 W-Boson Helicity in Top-Quark Decay:** The Standard Model dictates that the top quark has the same vector-minus-axial-vector ( $V - A$ ) charged-current weak interactions  $\left(-i\frac{g}{\sqrt{2}}V_{tb}\gamma^\mu\frac{1}{2}(1-\gamma_5)\right)$  as all the other fermions. In the SM, the fraction of top-quark decays to longitudinally polarized  $W$  bosons is similar to its Yukawa coupling and hence enhanced with respect to the weak coupling. It is expected to be [179]  $\mathcal{F}_0^{\text{SM}} \approx x/(1+x)$ ,  $x = m_t^2/2M_W^2$  ( $\mathcal{F}_0^{\text{SM}} \sim 70\%$  for  $m_t = 175 \text{ GeV}/c^2$ ). Fractions of left-handed, right-handed, or longitudinal  $W$  bosons are denoted as  $\mathcal{F}_-$ ,  $\mathcal{F}_+$ , and  $\mathcal{F}_0$  respectively. In the SM,  $\mathcal{F}_-$  is expected to be  $\approx 30\%$  and  $\mathcal{F}_+ \approx 0\%$ . Predictions for the  $W$  polarization fractions at NNLO in QCD are available [180].

The Tevatron and the LHC experiments use various techniques to measure the helicity of the  $W$  boson in top-quark decays, in both the lepton+jets and in dilepton channels in  $t\bar{t}$  production.

The first method uses a kinematic fit, similar to that used in the lepton+jets mass analyses, but with the top-quark mass constrained to a fixed value, to improve the reconstruction of final-state observables, and render the under-constrained dilepton channel solvable. Alternatively, in the dilepton channel the final-state momenta can also be obtained through an algebraic solution of the kinematics. The distribution of the helicity angle ( $\cos \theta^*$ ) between the lepton and the  $b$  quark in the  $W$  rest frame provides the most direct measure of the  $W$  helicity. In a simplified version of this approach, the  $\cos \theta^*$  distribution is reduced to a forward-backward asymmetry.

The second method ( $p_T^\ell$ ) uses the different lepton  $p_T$  spectra from longitudinally or transversely polarized  $W$ -decays to determine the relative contributions.

A third method uses the invariant mass of the lepton and the  $b$ -quark in top-quark decays ( $M_{\ell b}^2$ ) as an observable, which is directly related to  $\cos \theta^*$ .

At the LHC, top-quark pairs in the dilepton channels are reconstructed by solving a set of six independent kinematic equations on the missing transverse energy in  $x$ - and in  $y$ -direction, two  $W$ -masses, and the two top/antitop-quark masses. In addition, the two jets with the largest  $p_T$  in the event are interpreted as  $b$ -jets. The pairing of the jets to the charged leptons is based on the minimization of the sum of invariant masses  $M_{min}$ . Simulations show that this criterion gives the correct pairing in 68% of the events.

Finally, the Matrix Element method (ME) has also been used, in which a likelihood is formed from a product of event probabilities calculated from the ME for a given set of measured kinematic variables and assumed  $W$ -helicity fractions. The results of recent CDF, DØ, ATLAS, and CMS analyses are summarized in Table 2.

**Table 2:** Measurement and 95% C.L. upper limits of the  $W$  helicity in top-quark decays. The table includes both preliminary, as of August 2015, and published results. A full set of published results is given in the Listings.

$W$ Helicity	Source	$\int \mathcal{L} dt$ (fb $^{-1}$ )	Ref.	Method
$\mathcal{F}_0 = 0.722 \pm 0.081$	CDF+DØ Run II	2.7-5.4	[181]	$\cos \theta^*$ 2-par.
$\mathcal{F}_0 = 0.682 \pm 0.057$	CDF+DØ Run II	2.7-5.4	[181]	$\cos \theta^*$ 1-par.
$\mathcal{F}_0 = 0.726 \pm 0.094$	CDF Run II	8.7	[182]	ME 2-param
$\mathcal{F}_0 = 0.67 \pm 0.07$	ATLAS (7 TeV)	1.0	[183]	$\cos \theta^*$ 3-par.
$\mathcal{F}_0 = 0.682 \pm 0.045$	CMS (7 TeV)	5.0	[184]	$\cos \theta^*$ 3-par.
$\mathcal{F}_0 = 0.626 \pm 0.059$	ATLAS+CMS (7 TeV)	2.2	[185]	$\cos \theta^*$ 3-par.
$\mathcal{F}_0 = 0.659 \pm 0.027$	CMS (8 TeV)	19.6	[186]	$\cos \theta^*$ 3-par.
$\mathcal{F}_0 = 0.720 \pm 0.054$	CMS (8 TeV)	19.7	[187]	$\cos \theta^*$ 3-par.
$\mathcal{F}_0 = 0.653 \pm 0.029$	CMS (8 TeV)	19.7	[188]	$\cos \theta^*$ 3-par.
$\mathcal{F}_+ = -0.033 \pm 0.046$	CDF+DØ Run II	2.7-5.4	[181]	$\cos \theta^*$ 2-par.
$\mathcal{F}_+ = -0.015 \pm 0.035$	CDF+DØ Run II	2.7-5.4	[181]	$\cos \theta^*$ 1-par.
$\mathcal{F}_+ = -0.045 \pm 0.073$	CDF Run II	8.7	[182]	ME 2-par.
$\mathcal{F}_+ = 0.01 \pm 0.05$	ATLAS (7 TeV)	1.0	[183]	$\cos \theta^*$ 3-par.
$\mathcal{F}_+ = 0.008 \pm 0.018$	CMS (7 TeV)	5.0	[184]	$\cos \theta^*$ 3-par.
$\mathcal{F}_+ = 0.015 \pm 0.034$	ATLAS+CMS (7 TeV)	2.2	[185]	$\cos \theta^*$ 3-par.
$\mathcal{F}_+ = -0.009 \pm 0.021$	CMS (8 TeV)	19.6	[186]	$\cos \theta^*$ 3-par.
$\mathcal{F}_+ = -0.018 \pm 0.022$	CMS (8 TeV)	19.7	[187]	$\cos \theta^*$ 3-par.
$\mathcal{F}_+ = 0.018 \pm 0.027$	CMS (8 TeV)	19.7	[188]	$\cos \theta^*$ 3-par.

The datasets are now large enough to allow for a simultaneous fit of  $\mathcal{F}_0$ ,  $\mathcal{F}_-$  and  $\mathcal{F}_+$ , which we denote by ‘3-param’ or  $\mathcal{F}_0$  and  $\mathcal{F}_+$ , which we denote by ‘2-param’ in the table. Results with either  $\mathcal{F}_0$  or  $\mathcal{F}_+$  fixed at its SM value are denoted ‘1-param’. For the simultaneous fits, the correlation coefficient between the two values is about  $-0.8$ . A complete set of published results can be found in the Listings. All results are in agreement with the SM expectation.

CDF and DØ combined their results based on  $2.7 - 5.4$  fb $^{-1}$  [181] for a top-quark mass of  $172.5$  GeV/ $c^2$ . ATLAS presents results from  $1.04$  fb $^{-1}$  of  $\sqrt{s} = 7$  TeV data using a template method for the  $\cos \theta^*$  distribution and angular asymmetries from the unfolded  $\cos \theta^*$  distribution in the lepton+jets and the dilepton channel [183]. CMS performs a similar measurement based on template fits to the  $\cos \theta^*$  distribution with  $5.0$  fb $^{-1}$  of  $7$  TeV data in the lepton+jets final state [184]. As the polarization of the  $W$  bosons in top-quark decays is sensitive to the  $Wtb$  vertex Lorentz structure and anomalous couplings, both experiments also derive limits on anomalous contributions to the  $Wtb$  couplings. Recently, both experiments also combined their results from  $7$  TeV data to obtain values on the helicity fractions as well as limits on anomalous couplings [185].

CMS came out with a measurement of the  $W$ -helicity fractions in  $19.6$  fb $^{-1}$  of muon+jets events recorded at  $8$  TeV [186]. Also, using the same dataset a first measurement of the  $W$ -boson helicity in top-quark decays was made in electroweak single top production [187], yielding similarly precise and consistent results.

**C.2.4 Top-Quark Electroweak Charges:** The top quark is the only quark whose electric charge has not been measured

through production at threshold in  $e^+e^-$  collisions. Furthermore, it is the only quark whose electromagnetic coupling has not been observed and studied until recently. Since the CDF and DØ analyses on top-quark production did not associate the  $b$ ,  $\bar{b}$ , and  $W^\pm$  uniquely to the top or antitop, decays such as  $t \rightarrow W^+\bar{b}, \bar{t} \rightarrow W^-b$  were not excluded. A charge  $4/3$  quark of this kind is consistent with current electroweak precision data. The  $Z \rightarrow \ell^+\ell^-$  and  $Z \rightarrow b\bar{b}$  data, in particular the discrepancy between  $A_{LR}$  from SLC at SLAC and  $A_{FB}^{0,b}$  of  $b$ -quarks and  $A_{FB}^{0,\ell}$  of leptons from LEP at CERN, can be fitted with a top quark of mass  $m_t = 270$  GeV/ $c^2$ , provided that the right-handed  $b$  quark mixes with the isospin  $+1/2$  component of an exotic doublet of charge  $-1/3$  and  $-4/3$  quarks,  $(Q_1, Q_4)_R$  [189,190].

DØ studies the top-quark charge in double-tagged lepton+jets events, CDF does it in single tagged lepton+jets and dilepton events. Assuming the top- and antitop-quarks have equal but opposite electric charge, then reconstructing the charge of the  $b$ -quark through jet charge discrimination techniques, the  $|Q_{top}| = 4/3$  and  $|Q_{top}| = 2/3$  scenarios can be differentiated. For the exotic model of Chang *et al.* [190] with a top-quark charge  $|Q_{top}| = 4/3$ , DØ excludes the exotic model at 91.2% C.L. [191] using  $370$  pb $^{-1}$ , while CDF excludes the model at 99% C.L. [192] in  $5.6$  fb $^{-1}$ . Recently, DØ excluded the model at a significance greater than 5 standard deviations using  $5.3$  fb $^{-1}$  and set an upper limit of  $0.46$  on the fraction of such quarks in the selected sample [193]. All those results indicate that the observed particle is indeed consistent with being a SM  $|Q_{top}| = 2/3$  quark.

In  $2.05$  fb $^{-1}$  at  $\sqrt{s} = 7$  TeV, ATLAS performed a similar analysis, reconstructing the  $b$ -quark charge either via a jet-charge technique or via the lepton charge in soft muon decays in combination with a kinematic likelihood fit. They measure the top-quark charge to be  $0.64 \pm 0.02(stat.) \pm 0.08(syst.)e$  from the charges of the top-quark decay products in single lepton  $t\bar{t}$  events, and hence exclude the exotic scenario with charge  $-4/3$  at more than  $8\sigma$  [194].

In  $4.6$  fb $^{-1}$  at  $\sqrt{s} = 7$  TeV, CMS discriminates between the Standard Model and the exotic top-quark charge scenario in the muon+jets final states in  $t\bar{t}$  events. They exploit the charge correlation between high- $p_t$  muons from  $W$ -boson decays and soft muons from  $B$ -hadron decays in  $b$ -jets. Using an asymmetry technique, where  $A = -1$  represent the exotic  $q = -4/3$  scenario and  $A = +1$  the Standard Model  $q = +2/3$  scenario, they find  $A_{meas} = 0.97 \pm 0.12(stat.) \pm 0.31(sys.)$ , which agrees with the Standard Model expectation and excludes the exotic scenario at 99.9% C.L. [195].

The electromagnetic or the weak coupling of the top quark can be probed directly by investigating  $t\bar{t}$  events with an additional gauge boson, like  $t\bar{t}\gamma$  and  $t\bar{t}Z$  events.

CDF performs a search for events containing a lepton, a photon, significant missing transverse momentum, and a jet identified as containing a  $b$ -quark and at least three jets and large total transverse energy in  $6.0$  fb $^{-1}$ . They reported evidence for the observation of  $t\bar{t}\gamma$  production with a cross

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section  $\sigma_{t\bar{t}\gamma} = 0.18 \pm 0.08$  pb and a ratio of  $\sigma_{t\bar{t}\gamma}/\sigma_{t\bar{t}} = 0.024 \pm 0.009$  [196].

ATLAS performed a first measurement of the  $t\bar{t}\gamma$  cross section in pp collisions at  $\sqrt{s} = 7$  TeV using  $4.6 \text{ fb}^{-1}$  of data. Events are selected that contain a large transverse momentum electron or muon and a large transverse momentum photon, yielding 140 and 222 events in the electron and muon samples, respectively. The production of  $t\bar{t}\gamma$  events is observed with a significance of 5.3% standard deviations. The resulting cross section times branching ratio into the single lepton channel for  $t\bar{t}\gamma$  production with a photon with transverse momentum above 20 GeV is  $\sigma^{\text{fid.}}(t\bar{t}\gamma) \times Br = 63 \pm 8(\text{stat.})_{-13}^{+17}(\text{syst.}) \pm 1(\text{lumi.})$  pb per lepton flavour [198], which is consistent with leading-order theoretical calculations. Using  $19.7 \text{ fb}^{-1}$  of data at 8 TeV, CMS performs a similar measurement of the  $t\bar{t}\gamma$  production cross section in the muon+jets decay mode with a photon transverse momentum above 20 GeV and a separation  $\Delta R(\gamma, b/\bar{b}) > 0.1$ . They obtain a normalized cross section  $R = \sigma_{t\bar{t}+\gamma}/\sigma_{t\bar{t}} = (1.07 \pm 0.07(\text{stat.}) \pm 0.27(\text{syst.})) \times 10^{-2}$  and a cross section  $\sigma_{t\bar{t}+\gamma} = 2.4 \pm 0.2(\text{stat.}) \pm 0.6(\text{syst.})$  pb [199], consistent with the Standard Model expectations. A real test, however, of the vector and axial vector couplings in  $t\bar{t}\gamma$  events or searches for possible tensor couplings of top-quarks to photons will only be feasible with an integrated luminosity of several hundred  $\text{fb}^{-1}$  in the future.

ATLAS and CMS also studied the associate production of top-antitop quark pairs along with an electroweak gauge boson, where in the Standard Model the  $W$ -boson is expected to be produced via initial state radiation, while the  $Z$ -boson can also be radiated from a final-state top-quark and hence provides sensitivity to the top-quark neutral current weak gauge coupling, which implies a sensitivity to the third component of the top-quark's weak isospin.

CMS performed measurements of the  $t\bar{t}W$  and  $t\bar{t}Z$  production cross section at  $\sqrt{s} = 7$  TeV with  $5 \text{ fb}^{-1}$ , yielding results at about 3 standard deviations significance [200]. ATLAS performed a similar analysis with  $4.7 \text{ fb}^{-1}$  in the three-lepton channel and set an upper limit of 0.71 pb at 95% C.L. [201].

Using  $20.3 \text{ fb}^{-1}$  of 8 TeV data, ATLAS performs a simultaneous measurement of the  $t\bar{t}W$  and  $t\bar{t}Z$  cross section. They observe the  $t\bar{t}W$  and  $t\bar{t}Z$  production at the  $5.0\sigma$  and  $4.2\sigma$  level, respectively, yielding  $\sigma_{t\bar{t}W} = 369_{-91}^{+100}$  fb and  $\sigma_{t\bar{t}Z} = 176_{-52}^{+58}$  fb [202]. CMS performs an analysis where signal events are identified by matching reconstructed objects in the detector to specific final state particles from  $t\bar{t}W$  and  $t\bar{t}Z$  decays. using  $19.5 \text{ fb}^{-1}$  of 8 TeV data. They obtain  $\sigma_{t\bar{t}W} = 382_{-102}^{+117}$  fb and  $\sigma_{t\bar{t}Z} = 242_{-55}^{+65}$  fb, yielding a significance of 4.8 and 6.4 standard, respectively [203]. These measurements are used to set bounds on five anomalous dimension-six operators that would affect the  $t\bar{t}W$  and  $t\bar{t}Z$  cross sections.

## C.3 Searches for Physics Beyond the Standard Model

The top quark plays a special role in the SM. Being the only quark with a coupling to the Higgs boson of order one,

it provides the most important contributions to the quadratic radiative corrections to the Higgs mass raising the question of the naturalness of the SM. It is therefore very common for models where the naturalness problem is addressed to have new physics associated with the top quark. In SUSY, for instance, naturalness predicts the scalar top partners to be the lightest among the squarks and to be accessible at the LHC energies (see the review "Supersymmetry: Theory"). In models where the Higgs is a pseudo-Goldstone boson, such as Little Higgs models, naturalness predicts the existence of partners of the top quarks with the same spin and color, but with different electroweak couplings, the so-called vectorial  $t'$ . Stops and  $t'$ 's are expected to have sizable branching ratios to top quarks. Another intriguing prediction of SUSY models with universal couplings at the unification scale is that for a top-quark mass close to the measured value, the running of the Yukawa coupling down to 1 TeV naturally leads to the radiative breaking of the electroweak symmetry [204]. In fact, the top quark plays a role in the dynamics of electroweak symmetry breaking in many models. One example is topcolor [205], where a large top-quark mass can be generated through the formation of a dynamic  $t\bar{t}$  condensate,  $X$ , which is formed by a new strong gauge force coupling preferentially to the third generation. Another example is topcolor-assisted technicolor [206], predicting the existence of a heavy  $Z'$  boson that couples preferentially to the third generation of quarks. If light enough such a state might be directly accessible at the present hadron collider energies, or if too heavy, lead to four-top interactions possibly visible in the  $t\bar{t}t\bar{t}$  final state, for which limits on production cross sections at the LHC  $\sqrt{s} = 8$  TeV exist [207,208].

Current strategies to search for new physics in top-quark events at hadron colliders are either tailored to the discovery of specific models or model independent. They can be broadly divided in two classes. In the first class new resonant states are looked for through decay processes involving the top quarks. Current searches for bosonic resonances in  $t\bar{t}$  final states, or for direct stop and  $t'$  production, or for a charged Higgs in  $H^+ \rightarrow t\bar{b}$  fall in the category. On the other hand, if new states are too heavy to be directly produced, they might still give rise to deviations from the SM predictions for the strength and Lorentz form of the top-quark couplings to other SM particles. Accurate predictions and measurements are therefore needed and the results be efficiently systematized in the framework of an effective field theory [210,211]. For instance, the on-going efforts to constrain the structure of the top couplings to vector bosons ( $g, \gamma, Z, W$ ) and to the Higgs boson, including flavor-changing neutral currents involving the top quark [212], fall in this second category.

**C.3.1 New Physics in Top-Quark Production:** Theoretical [213–215] and experimental efforts have been devoted to the searches of  $t\bar{t}$  resonances.

At the Tevatron, both the CDF and DØ collaborations have searched for resonant production of  $t\bar{t}$  pairs in the lepton+jets

channel [216,217]. In both analyses, the data indicate no evidence of resonant production of  $t\bar{t}$  pairs. They place upper limits on the production cross section times branching fraction to  $t\bar{t}$  in comparison to the prediction for a narrow ( $\Gamma_{Z'} = 0.012M_{Z'}$ ) leptophobic topcolor  $Z'$  boson. Within this model, they exclude  $Z'$  bosons with masses below 915 (CDF-full data set) and 835 ( $D\bar{O}$ ,  $5\text{ fb}^{-1}$ )  $\text{GeV}/c^2$  at the 95% C.L. These limits turn out to be independent of couplings of the  $t\bar{t}$  resonance (pure vector, pure axial-vector, or SM-like  $Z'$ ). A similar analysis has been performed by CDF in the all-jets channel using  $2.8\text{ fb}^{-1}$  of data [218].

At the LHC, both the CMS and ATLAS collaborations have searched for resonant production of  $t\bar{t}$  pairs, employing different techniques and final-state signatures (all-jets, lepton+jets, dilepton) at  $\sqrt{s} = 7$  and 8 TeV. In the low mass range, from the  $t\bar{t}$  threshold to about one TeV, standard techniques based on the reconstruction of each of the decay objects (lepton, jets and  $b$ -jets, missing  $E_T$ ) are used to identify the top quarks, while at higher invariant mass, the top quarks are boosted and the decay products more collimated and can appear as large-radius jets with substructure. Dedicated reconstruction techniques have been developed in recent years for boosted top quarks [219] that are currently employed at the LHC. Most of the analyses are model-independent (i.e., no assumption on the quantum numbers of the resonance is made) yet they assume a small width and no signal-background interference.

Using dilepton and lepton+jets signatures in a data set corresponding to an integrated luminosity of  $5.0\text{ fb}^{-1}$ , the CMS collaboration finds no significant deviations from the SM background. In the dilepton analysis, upper limits are presented for the production cross section times branching fraction of top quark-antiquark resonances for masses from 750 to 3000  $\text{GeV}/c^2$ . In particular, the existence of a leptophobic topcolor particle  $Z'$  is excluded at the 95% confidence level for resonance masses  $M_{Z'} < 1.3$  (1.9)  $\text{TeV}/c^2$  for  $\Gamma_{Z'} = 0.012(0.1)M_{Z'}$  [220]. Using a lepton+jets sample, results are obtained from the combination of two dedicated searches optimized for boosted production and production at threshold. In this case, topcolor  $Z'$  bosons with narrow (wide) width are excluded at 95% confidence level for masses below 1.49 (2.04)  $\text{TeV}/c^2$  and an upper limit of 0.3 (1.3) pb or lower is set on the production cross section times branching fraction for resonance masses above 1  $\text{TeV}/c^2$ . Kaluza-Klein excitations of a gluon with masses below 1.82  $\text{TeV}/c^2$  (at 95% confidence level) in the Randall-Sundrum model are also excluded, and an upper limit of 0.7 pb or lower is set on the production cross section times branching fraction for resonance masses above 1  $\text{TeV}/c^2$  [221]. In  $19.7\text{ fb}^{-1}$  of 8 TeV data, CMS recently updated their measurement in the lepton+jets and the all-jets channel to obtain an exclusion of  $M_{Z'} < 2.1(2.7)\text{ TeV}/c^2$  for  $\Gamma_{Z'} = 0.013(0.1)M_{Z'}$  and gluon masses below 2.5  $\text{TeV}/c^2$  in Randall-Sundrum models at 95% C.L. [222]. These limits have been improved in a recent analysis which uses events with three different final states, defined by the number of leptons and optimized for reconstruction

of top quarks with high Lorentz boosts [223]. For example, in this analysis a narrow leptophobic topcolor  $Z'$  resonance with a mass below 2.4 TeV is excluded at 95% confidence level.

The ATLAS collaboration has performed a search for resonant  $t\bar{t}$  production in the lepton+jets channel using  $4.7\text{ fb}^{-1}$  ( $19.7\text{ fb}^{-1}$ ) of proton-proton (pp) collision data collected at a center-of-mass energy  $\sqrt{s} = 7(8)\text{ TeV}$  [224,225]. The  $t\bar{t}$  system is reconstructed using both small-radius and large-radius jets, the latter being supplemented by a jet substructure analysis. A search for local excesses in the number of data events compared to the Standard Model expectation in the  $t\bar{t}$  invariant mass spectrum is performed. No evidence for a  $t\bar{t}$  resonance is found and 95% confidence-level limits on the production rate are determined for massive states predicted in two benchmark models. The most stringent limits come from the sample collected at 8 TeV. The upper limits on the cross section times branching ratio of a narrow  $Z'$  boson decaying to top-quark pairs range from 4.2 pb for a resonance mass of 0.4  $\text{TeV}/c^2$  to 0.03 pb for a mass of 3  $\text{TeV}/c^2$ . A narrow leptophobic topcolor  $Z'$  boson with a mass below 1.8  $\text{TeV}/c^2$  is excluded. Upper limits are set on the cross section times branching ratio for a broad color-octet resonance with  $\Gamma/m = 15\%$  decaying to  $t\bar{t}$ . These range from 2.5 pb for a mass of 0.4  $\text{TeV}/c^2$  to 0.03 pb for a mass of 3  $\text{TeV}/c^2$ . A Kaluza-Klein excitation of the gluon in a Randall-Sundrum model (a slightly different model is used compared to CMS) is excluded for masses below 2.2  $\text{TeV}/c^2$ .

ATLAS has also conducted a search in the all-jet final state at 7 TeV corresponding to an integrated luminosity of  $4.7\text{ fb}^{-1}$  [226]. The  $t\bar{t}$  events are reconstructed by selecting two top quarks in their fully hadronic decay modes which are reconstructed using the Cambridge/Aachen jet finder algorithm with a radius parameter of 1.5. The substructure of the jets is analysed using the HEPTopTagger algorithm [227] to separate top-quark jets from those originating from gluons and lighter quark jets. The invariant mass spectrum of the data is compared to the SM prediction, and no evidence for resonant production of top-quark pairs is found. The data are used to set upper limits on the cross section times branching ratio for resonant  $t\bar{t}$  production in two models at 95% confidence level. Leptophobic  $Z'$  bosons with masses between 700 and 1000  $\text{GeV}/c^2$  as well as 1280 – 1320  $\text{GeV}/c^2$  and Kaluza-Klein-Gluons with masses between 700 and 1620  $\text{GeV}/c^2$  are excluded at the 95% confidence level.

Heavy charged bosons, such as  $W'$  or  $H^+$ , can also be searched for in  $t\bar{b}$  final states (for more information see the review "W'-boson searches" and "Higgs Bosons: theory and searches"). Other resonances are searched for in final states such as  $tZ, t\bar{j}, tH, tW, bW$ .

For instance, ATLAS has performed a search for  $t$ -jet resonances in the lepton+jets channel of  $t\bar{t}$ +jets events in  $4.7\text{ fb}^{-1}$  at  $\sqrt{s} = 7\text{ TeV}$  [228]. A heavy new particle, assumed to be produced singly in association with a  $t(\bar{t})$  quark, decays to a  $t(\bar{t})$  quark and a light flavor quark, leading to a color singlet (triplet) resonance in the  $t(\bar{t})$ +jet system. The full 2011 ATLAS

# Quark Particle Listings

## *t*

$pp$  collision dataset from the LHC ( $4.7 \text{ fb}^{-1}$ ) is used to select  $t\bar{t}$  events. The data are consistent with the SM expectation and a new particle with mass below 350 (430)  $\text{GeV}/c^2$  for  $W$  (color triplet) models is excluded with a 95% confidence level, assuming unit right-handed coupling. ATLAS has conducted a search for the single and pair production of a new charge  $+2/3$  quark ( $T$ ) decaying via  $T \rightarrow Zt$  (and also  $-1/3$  quark ( $B$ ) decaying via  $B \rightarrow Zb$ ) in a dataset corresponding to  $20.3 \text{ fb}^{-1}$  luminosity at  $\sqrt{s} = 8 \text{ TeV}$  [229]. Selected events contain a high transverse momentum  $Z$ -boson candidate reconstructed from a pair of oppositely charged electrons or muons. Additionally, the presence of at least two jets possessing properties consistent with the decay of a  $b$ -hadron is required, as well as large total transverse momentum of all central jets in the event. No significant excess of events above the SM expectation is observed, and upper limits are derived for vector-like quarks of various masses in a two-dimensional plane of branching ratios. Under branching ratio assumptions corresponding to a weak-isospin singlet scenario, a  $T$  quark with mass lower than  $655 \text{ GeV}/c^2$  is excluded at the 95% confidence level. Under branching ratio assumptions corresponding to a particular weak-isospin doublet scenario, a  $T$  quark with mass lower than  $735 \text{ GeV}/c^2$  is excluded at the 95% confidence level.

A complementary search [208] in the lepton+jets final state of the same dataset, characterized by an isolated electron or muon with moderately high transverse momentum, significant missing transverse momentum, and multiple jets is performed to look for  $T(B) \rightarrow Wb, Zt, Ht(Wt, Zb, Hb)$ , decays. No significant excess of events above the SM expectation is observed, and upper limits are derived for vector-like quarks of various masses under several branching ratio hypotheses. The 95% C.L. observed lower limits on the  $T$  quark mass range between 715  $\text{GeV}$  and 950  $\text{GeV}$  for all possible values of the branching ratios into the three decay modes. In addition this study provides limits on four top-quark production and production of two positively-charged top quarks. No significant excess of events over the background expectation is observed. The four top-quark production cross section must be less than 23  $\text{fb}$  in the SM and less than 12  $\text{fb}$  for production via a contact interaction; in the case of sgluon pair production decaying to  $t\bar{t}$ , where a sgluon is a scalar partner of the gluino [209], the mass of a sgluon must be greater than  $1.06 \text{ TeV}/c^2$ . Finally, limits in the context of models featuring two extra dimensions are also set.

In many models top-quark partners preferably decay to top quarks and weakly interacting neutral stable particles, i.e., possibly dark matter candidates, that are not detected. An observable especially sensitive to new physics effects in  $t\bar{t}$  production is therefore the missing momentum.

CMS has presented a differential cross section measurement of top-quark pair production with missing transverse energy using  $20 \text{ fb}^{-1}$  at 8  $\text{TeV}$  [230]. The results are consistent with the predictions of the SM. More recently, CMS has presented a search for particle dark matter produced in association with a pair of top quarks in  $19.7 \text{ fb}^{-1}$  of data at 8  $\text{TeV}$  [231].

This search requires the presence of one lepton, multiple jets, and large missing transverse energy. No excess of events is found above the SM expectation, and upper limits are derived on the production cross section. Cross sections larger than 20 to 55  $\text{fb}$  are excluded at 90% C.L. for dark matter particles with the masses ranging from 1 to 1000  $\text{GeV}$ . Interpreting the findings in the context of a scalar contact interaction between fermionic dark matter particles and top quarks, lower limits on the interaction scale are set. Assuming a dark matter particle with a mass of 100  $\text{GeV}$ , values of the interaction scale below 118  $\text{GeV}$  are excluded at 90% C.L. An analogous search, at a center-of-mass energy of 7  $\text{TeV}$  in  $1.04 \text{ fb}^{-1}$  of data has been performed by ATLAS [232]. The search is carried out in the lepton+jets channel. The results are interpreted in terms of a model where new top-quark partners are pair-produced and each decay to an on-shell top (or antitop) quark and a long-lived undetected neutral particle. The data are found to be consistent with SM expectations. A limit at 95% C.L. is set excluding a cross-section times branching ratio of 1.1  $\text{pb}$  for a top-partner mass of  $420 \text{ GeV}/c^2$  and a neutral particle mass less than 10  $\text{GeV}/c^2$ . In a model of exotic fourth generation quarks, top-partner masses are excluded up to  $420 \text{ GeV}/c^2$  and neutral particle masses up to  $140 \text{ GeV}/c^2$ .

Flavor-changing-neutral-currents (FCNC) are hugely suppressed in the SM, and non zero only due to the large mass hierarchy between the top quark and the other quarks. Several observables are accessible at colliders to test and constrain such couplings.

CMS has performed several studies on the search for FCNC in top-quark production. They have considered single top quark production in the  $t$ -channel in  $5 \text{ fb}^{-1}$  integrated luminosity at 7  $\text{TeV}$  [233]. Events with the top quark decaying into a muon, neutrino and  $b$ -quark are selected. The upper limits on effective coupling strength can be translated to the 95% upper limits on the corresponding branching ratios  $B(t \rightarrow gu) \leq 3.55 \cdot 10^{-4}$ ,  $B(t \rightarrow gc) \leq 3.44 \cdot 10^{-3}$ . They have performed a search for a single top quark produced in association with a photon in  $19.1 \text{ fb}^{-1}$  integrated luminosity at 8  $\text{TeV}$  [234]. The event selection requires the presence of one isolated muon and jets in the final state. The upper limits on effective coupling strength can be translated to the 95% upper limits on the corresponding branching ratios  $B(t \rightarrow \gamma u) \leq 0.0161\%$ ,  $B(t \rightarrow \gamma c) \leq 0.182\%$ .

ATLAS has presented results on the search for single top-quark production via FCNC's in strong interactions using data collected at  $\sqrt{s}=8 \text{ TeV}$  and corresponding to an integrated luminosity of  $20.3 \text{ fb}^{-1}$ . Flavor-changing-neutral-current events are searched for in which a light quark ( $u$  or  $c$ ) interacts with a gluon to produce a single top quark, either with or without the associated production of another light quark or gluon. Candidate events of top quarks decaying into leptons and jets are selected and classified into signal- and background-like events using a neural network. The observed 95% C.L. limit is  $\sigma_{qq \rightarrow t} \times B(t \rightarrow Wb) < 3.4 \text{ pb}$  that can be interpreted as limits on the branching ratios,  $B(t \rightarrow ug) < 4 \cdot 10^{-5}$  and  $B(t \rightarrow cg) <$

$1.7 \cdot 10^{-4}$  [235]. This result supersedes the corresponding 7 TeV analysis in  $2 \text{ fb}^{-1}$  [236].

Constraints on FCNC couplings of the top quark can also be obtained from searches for anomalous single top-quark production in  $e^+e^-$  collisions, via the process  $e^+e^- \rightarrow \gamma, Z^* \rightarrow t\bar{q}$  and its charge-conjugate ( $q = u, c$ ), or in  $e^\pm p$  collisions, via the process  $e^\pm u \rightarrow e^\pm t$ . For a leptonic  $W$  decay, the topology is at least a high- $p_T$  lepton, a high- $p_T$  jet and missing  $E_T$ , while for a hadronic  $W$ -decay, the topology is three high- $p_T$  jets. Limits on the cross section for this reaction have been obtained by the LEP collaborations [237] in  $e^+e^-$  collisions, and by H1 [238] and ZEUS [239] in  $e^\pm p$  collisions. When interpreted in terms of branching ratios in top decay [240,241], the LEP limits lead to typical 95% C.L. upper bounds of  $B(t \rightarrow qZ) < 0.137$ . Assuming no coupling to the  $Z$  boson, the 95% C.L. limits on the anomalous FCNC coupling  $\kappa_\gamma < 0.13$  and  $< 0.27$  by ZEUS and H1, respectively, are stronger than the CDF limit of  $\kappa_\gamma < 0.42$ , and improve over LEP sensitivity in that domain. The H1 limit is slightly weaker than the ZEUS limit due to an observed excess of five-candidate events over an expected background of  $3.2 \pm 0.4$ . If this excess is attributed to FCNC top-quark production, this leads to a total cross section of  $\sigma(ep \rightarrow e + t + X, \sqrt{s} = 319 \text{ GeV}) < 0.25 \text{ pb}$  [238,242].

**C.3.2 New Physics in Top-Quark decays:** The large sample of top quarks produced at the Tevatron and the LHC allows to measure or set stringent limits on the branching ratios of rare top-quark decays. For example, the existence of a light  $H^+$  can be constrained by looking for  $t \rightarrow H^+ b$  decay, in particular with tau-leptons in the final state (for more information see the review "Higgs Bosons: theory and searches").

A first class of searches for new physics focuses on the structure of the  $Wtb$  vertex. Using up to  $2.7 \text{ fb}^{-1}$  of data, DØ has measured the  $Wtb$  coupling form factors by combining information from the  $W$ -boson helicity in top-quark decays in  $t\bar{t}$  events and single top-quark production, allowing to place limits on the left-handed and right-handed vector and tensor couplings [243–245].

ATLAS has published the results of a search for  $CP$  violation in the decay of single top quarks produced in the  $t$ -channel where the top quarks are predicted to be highly polarized, using the lepton+jets final state [246]. The data analyzed are from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  and correspond to an integrated luminosity of  $4.7 \text{ fb}^{-1}$ . In the Standard Model, the couplings at the  $Wtb$  vertex are left-handed, right-handed couplings being absent. A forward-backward asymmetry with respect to the normal to the plane defined by the  $W$ -momentum and the top-quark polarization has been used to probe the complex phase of a possibly non-zero value of the right-handed coupling, signaling a source of  $CP$ -violation beyond the SM. The measured value of the asymmetry is  $0.031 \pm 0.065(stat.)^{+0.029}_{-0.031}(syst.)$  in good agreement with the Standard Model.

A second class of searches focuses on FCNC's in the top-quark decays. Both, CDF and DØ, have provided the first

limits for FCNC's in Run I and II. The most recent results from CDF give  $B(t \rightarrow qZ) < 3.7\%$  and  $B(t \rightarrow q\gamma) < 3.2\%$  at the 95% C.L. [247] while DØ [248,249] sets  $B(t \rightarrow qZ)(q = u, c \text{ quarks}) < 3.2\%$  at 95% C.L.,  $B(t \rightarrow gu) < 2.0 \cdot 10^{-4}$ , and  $B(t \rightarrow gc) < 3.9 \cdot 10^{-3}$  at the 95% C.L.

At the LHC, CMS has used a sample at a center-of-mass energy of 8 TeV corresponding to  $19.7 \text{ fb}^{-1}$  of integrated luminosity to perform a search for flavor changing neutral current top-quark decay  $t \rightarrow Zq$ . Events with a topology compatible with the decay chain  $t\bar{t} \rightarrow Wb + Zq \rightarrow \ell\nu b + \ell\ell q$  are searched for. There is no excess seen in the observed number of events relative to the SM prediction; thus no evidence for flavor changing neutral current in top-quark decays is found. A combination with a previous search at 7 TeV excludes a  $t \rightarrow Zq$  branching fraction greater than 0.05% at the 95% confidence level [250]. The ATLAS collaboration has also searched for FCNC processes in  $20.3 \text{ fb}^{-1}$  of  $t\bar{t}$  events with one top quark decaying through FCNC ( $t \rightarrow qZ$ ) and the other through the SM dominant mode ( $t \rightarrow bW$ ). Only the decays of the  $Z$  boson to charged leptons and leptonic  $W$  boson decays were considered as signal, leading to a final state topology characterized by the presence of three isolated leptons, at least two jets and missing transverse energy from the undetected neutrino. No evidence for an FCNC signal was found. An upper limit on the  $t \rightarrow qZ$  branching ratio of  $B(t \rightarrow qZ) < 7 \times 10^{-4}$  is set at the 95% confidence level [251], which supersedes previous results [252].

Another search for FCNCs is in the decay of a top-quark to a Higgs boson plus a light parton,  $t \rightarrow qH$ ,  $q = u, c$ . The CMS collaboration has performed two searches using a sample at a center-of-mass energy of 8 TeV corresponding to  $19.7 \text{ fb}^{-1}$  of integrated luminosity, one in a multi-lepton final state [253] and the other with the Higgs boson decaying to  $\gamma\gamma$  [254]. The first analysis sets an upper limit on the  $t \rightarrow cH$  branching ratio of  $B(t \rightarrow cH) < 0.93\%$  at 95% confidence level, while the second sets an upper limit on the  $t \rightarrow c(u)H$  branching ratios of  $B(t \rightarrow c(u)H) < 0.71(0.65)\%$  at 95% confidence level. The ATLAS collaboration considers  $t \rightarrow qH$ ,  $q = u, c$  with  $4.7 \text{ fb}^{-1}$  of  $t\bar{t}$  events at  $\sqrt{s} = 7 \text{ TeV}$  and  $20.3 \text{ fb}^{-1}$  of  $t\bar{t}$  events at  $\sqrt{s} = 8 \text{ TeV}$ . A combined measurement including  $H \rightarrow \gamma\gamma$  and  $H \rightarrow WW^*, \tau\tau$  modes yields a 95% C.L. upper limit of 0.46% and 0.45% on the branching ratios of  $B(t \rightarrow cH)$  and  $B(t \rightarrow uH)$ , respectively [255].

## D. Outlook

Top-quark physics at hadron colliders has developed into precision physics. Various properties of the top quark have been measured with high precision, where the LHC is about to or has already reached the precision of the Tevatron. Several  $\sqrt{s}$ -dependent physics quantities, such as the production cross-section, have been measured at several energies at the Tevatron and the LHC. Up to now, all measurements are consistent with the SM predictions and allow stringent tests of the underlying production mechanisms by strong and weak interactions. Given the very large event samples available at

# Quark Particle Listings

*t*

the LHC, top-quark properties will be further determined in  $t\bar{t}$  as well as in electroweak single top-quark production. At the Tevatron, the  $t$ – and  $s$ –channels for electroweak single top-quark production have been measured separately. At the LHC, significant progress has been achieved and all the three relevant channels are expected to be independently accessible in the near future. Furthermore,  $t\bar{t}\gamma$ ,  $t\bar{t}Z$ , and  $t\bar{t}W$  together with  $t\bar{t}H$  associated production will provide further information on the top-quark electroweak couplings. At the same time various models of physics beyond the SM involving top-quark production are being constrained. With the first results from LHC Run-II at a higher center-of-mass energy and much higher luminosity starting to be released, top-quark physics has the potential to shed light on open questions and new aspects of physics at the TeV scale.

## References

CDF note references can be retrieved from

[www-cdf.fnal.gov/physics/new/top/top.html](http://www-cdf.fnal.gov/physics/new/top/top.html),  
and DØ note references from [www-d0.fnal.gov/Run2Physics/WWW/documents/Run2Results.htm](http://www-d0.fnal.gov/Run2Physics/WWW/documents/Run2Results.htm),  
and ATLAS note references from <https://twiki.cern.ch/twiki/bin/view/AtlasPublic/TopPublicResults>,  
and CMS note references from <https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsTOP>.

1. M. Czakon, P. Fiedler, and A. Mitov, Phys. Rev. Lett. **110**, 252004 (2013).
2. M. Cacciari *et al.*, J. High Energy Phys.0809,127(2008); N. Kidonakis and R. Vogt, Phys. Rev. **D78**, 074005 (2008); S. Moch and P. Uwer, Nucl. Phys. (Proc. Supp.) **B183**, 75 (2008); S. Moch and P. Uwer, Phys. Rev. **D78**, 034003 (2008); U. Langenfeld, S. Moch, and P. Uwer, Phys. Rev. **D80**, 054009 (2009); M. Beneke *et al.* Phys. Lett. **B690**, 483 (2010); M. Beneke *et al.*, Nucl. Phys. **B855**, 695 (2012); V. Ahrens *et al.*, Phys. Lett. **B703**, 135 (2011); M. Cacciari *et al.*, Phys. Lett. **B710**, 612 (2012).
3. ATLAS, CMS, CDF, & DØ Collabs., [arXiv:1403.4427](https://arxiv.org/abs/1403.4427).
4. S. Cortese and R. Petronzio, Phys. Lett. **B253**, 494 (1991).
5. S. Willenbrock and D. Dicus, Phys. Rev. **D34**, 155 (1986).
6. N. Kidonakis, Phys. Rev. **D83**, 091503 (2011).
7. M. Brucherseifer, F. Caola, and K. Melnikov, Phys. Lett. **B736**, 58 (2014).
8. N. Kidonakis, Phys. Rev. **D81**, 054028 (2010).
9. N. Kidonakis, Phys. Rev. **D82**, 054018 (2010).
10. T. Tait and C.-P. Yuan. Phys. Rev. **D63**, 014018 (2001).
11. M. Jezabek and J.H. Kühn, Nucl. Phys. **B314**, 1 (1989).
12. I.I.Y. Bigi *et al.*, Phys. Lett. **B181**, 157 (1986).
13. A.H. Hoang *et al.*, Phys. Rev. **D65**, 014014 (2002).
14. K. Hagiwara, Y. Sumino, and H. Yokoya, Phys. Lett. **B666**, 71 (2008).
15. A. Czarnecki and K. Melnikov, Nucl. Phys. **B544**, 520 (1999); K.G. Chetyrkin *et al.*, Phys. Rev. **D60**, 114015 (1999).
16. S. Frixione, P. Nason, and B. Webber, J. High Energy Phys.08,007(2003); S. Frixione, P. Nason, and C. Oleari, J. High Energy Phys.07,070(2007); S. Frixione, P. Nason, and G. Ridolfi, J. High Energy Phys.07,126(2007) J.M. Campbell *et al.*, J. High Energy Phys.1504,114(2015).
17. S. Frixione *et al.*, J. High Energy Phys.06,092(2006); S. Frixione *et al.*, J. High Energy Phys.08,029(2008); S. Alioli *et al.*, J. High Energy Phys.09,111(2009); E. Re, Eur. Phys. J. **C71**, 1547 (2011); R. Frederix, E. Re, and P. Torrielli, J. High Energy Phys.12,130(2012).
18. S. Frixione and B.R. Webber, J. High Energy Phys.02,029(2002).
19. P. Nason, J. High Energy Phys.04,040(2004).
20. V.M. Abazov *et al.* (DØ Collab.), DØ-CONF-Note 6453 (2015).
21. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D88**, 091103 (2013).
22. T. Aaltonen *et al.* (CDF and DØ Collab.), Phys. Rev. **D89**, 072001 (2014).
23. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D89**, 091101 (2014).
24. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C74**, 3109 (2014).
25. ATLAS Collab., ATLAS-CONF-2011-121.
26. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.1205,059(2012).
27. ATLAS Collab., ATLAS-CONF-2011-140.
28. ATLAS Collab., ATLAS-CONF-2012-024.
29. ATLAS Collab., ATLAS-CONF-2012-031.
30. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C73**, 2328 (2013).
31. G. Aad *et al.* (ATLAS Collab.), Phys. Lett. **B717**, 89 (2012).
32. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.11,067(2012).
33. S. Chatrchyan *et al.* (CMS Collab.), Phys. Lett. **B720**, 83 (2013).
34. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.1305,065(2013).
35. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. **D85**, 112007 (2012).
36. S. Chatrchyan *et al.* (CMS Collab.), Eur. Phys. J. **C73**, 2386 (2013).
37. ATLAS & CMS Collabs., ATLAS-CONF-2012-134, CMS PAS TOP-12-003.
38. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D91**, 112013 (2015).
39. CMS Collab., CMS-PAS-TOP-12-006.
40. CMS Collab., CMS-PAS-TOP-13-004.
41. S. Chatrchyan *et al.* CMS Collab., [arxiv:1509.06076](https://arxiv.org/abs/1509.06076). Submitted to Eur. Phys. J. C.
42. ATLAS Collab., ATLAS-CONF-2014-053, CMS Collab., CMS-PAS-TOP-14-016.
43. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **115**, 112001 (2015).
44. ATLAS Collab., ATLAS-CONF-2015-033.



45. CMS Collab., CMS-PAS-TOP-15-005.
46. ATLAS Collab., ATLAS-CONF-2011-108.
47. V.M. Abazov *et al.* (DØ Collab.) Phys. Rev. Lett. **107**, 121802, (2011); D. Acosta *et al.* (CDF Collab.) Phys. Rev. Lett. **95**, 102002, (2005).
48. V. Khachatryan *et al.* (CMS Collab.), Phys. Lett. **B736**, 33 (2014).
49. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D67**, 012004 (2003).
50. T. Affolder *et al.* (CDF Collab.), Phys. Rev. **D64**, 032002 (2001).
51. M. Czakon, D. Heymes, A. Mitov, [arXiv:1511.00549](#).
52. T. Affolder *et al.* (CDF Collab.), Phys. Rev. Lett. **102**, 222003 (2009).
53. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D90**, 092006 (2014).
54. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C73**, 2261 (2013).
55. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D90**, 072004 (2014).
56. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys. **06**, 100(2015).
57. ATLAS Collab., ATLAS-CONF-2014-057 (2014).
58. CMS Collab., CMS-PAS-TOP-14-012.
59. S. Chatrchyan *et al.* (CMS Collab.), Eur. Phys. J. **C73**, 2339 (2013).
60. S. Chatrchyan *et al.* (CMS Collab.), Eur. Phys. J. **C75**, 542 (2015).
61. CMS Collab., CMS-PAS-TOP-15-010 (2015).
62. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C76**, 11 (2016).
63. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys. **01**, 020(2015).
64. S. Chatrchyan *et al.* (CMS Collab.), Eur. Phys. J. **C74**, 3014 (2014).
65. S. Chatrchyan *et al.* (CMS Collab.), Phys. Lett. **B746**, 132 (2015).
66. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D92**, 072005 (2015).
67. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **103**, 092001 (2009); V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D78**, 12005 (2008); V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **98**, 181802 (2007).
68. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **103**, 092002 (2009); T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D81**, 072003 (2010).
69. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D82**, 112005 (2010).
70. A. Heinson and T. Junk, Ann. Rev. Nucl. and Part. Sci. **61**, 171 (2011).
71. Tevatron Electroweak Working Group, [arXiv:0908.2171 \[hep-ex\]](#).
72. CDF Collab., CDF conference note 11113 (2014), DØ Collab., DØ conference note 6448 (2014).
73. T. Aaltonen *et al.* (CDF and DØ Collab.), Phys. Rev. Lett. **112**, 231803 (2014).
74. G. Aad *et al.*, ATLAS Collab., Phys. Rev. **D90**, 112006 (2014).
75. G. Aad *et al.*, ATLAS Collab., Phys. Lett. **B717**, 330 (2012).
76. S. Chatrchyan *et al.*, CMS Collab., J. High Energy Phys. **12**, 035(2012).
77. ATLAS Collab., ATLAS-CONF-2014-007.
78. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys. **06**, 090(2014).
79. CMS Collab., CMS-PAS-TOP-15-004.
80. C.D. White *et al.*, J. High Energy Phys. **11**, 74(2009).
81. S. Frixione *et al.*, J. High Energy Phys. **07**, 29(2008).
82. G. Aad *et al.* (ATLAS Collab.), Phys. Lett. **B716**, 142 (2012).
83. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **110**, 022003 (2012).
84. ATLAS Collab., ATLAS-CONF-2013-100.
85. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **112**, 231802 (2014).
86. ATLAS Collab., ATLAS-CONF-2014-052, CMS Collab., CMS-PAS-TOP-14-009.
87. ATLAS Collab., ATLAS-CONF-2011-118.
88. ATLAS Collab., ATLAS-CONF-2015-047.
89. CMS Collab., CMS-PAS-TOP-13-009.
90. CDF Collab., CDF conference note 10793 (2012).
91. CDF Collab., CDF conference note 11015 (2013).
92. V.M. Abazov *et al.* (DØ Collab.), Phys. Lett. **B726**, 656 (2013).
93. ATLAS Collab., ATLAS-CONF-2012-056.
94. S. Chatrchyan, *et al.* (CMS Collab.), Phys. Lett. **B720**, 83 (2013).
95. CMS Collab., CMS-PAS-TOP-12-011.
96. ATLAS and CMS Collab., ATLAS-CONF-2013-098, CMS-PAS-TOP-12-002.
97. CMS Collab., Phys. Rev. Lett. **112**, 231802 (2011).
98. CMS Collab., CMS-PAS-TOP-14-004.
99. CMS Collab., CMS-PAS-TOP-13-001.
100. M. Czakon, P. Fiedler, and A. Mitov Phys. Rev. Lett. **115**, 052001 (2015).
101. W. Hollik & D. Pagani Phys. Rev. **D84**, 093003 (2011).
102. W. Bernreuther & Z.G. Si, Phys. Rev. **D86**, 034026 (2012).
103. S. Jung, H. Murayama, A. Pierce, J.D. Wells, Phys. Rev. **D81**, 015004 (2010).
104. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **100**, 142002 (2008).
105. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **101**, 202001 (2008).
106. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D90**, 072011 (2014).
107. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D92**, 052007 (2015).
108. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D87**, 092002 (2013).
109. CDF Collab., CDF conference note 11161 (2015).
110. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys. **02**, 107(2014).
111. ATLAS Collab., [arXiv:1509.02358](#).
112. S. Chatrchyan *et al.* (CMS Collab.), Phys. Lett. **B717**, 129 (2012).

## Quark Particle Listings

 $t$ 

113. V. Khachatryan *et al.* (CMS Collab.), [arXiv:1508.03862](#), submitted to Phys. Rev. D.
114. V. Khachatryan *et al.* (CMS Collab.), [arXiv:1507.03119](#), submitted to Phys. Lett. B.
115. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D90**, 072001 (2014).
116. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D88**, 112002 (2013).
117. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D88**, 072003 (2013).
118. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.05,061(2015).
119. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.04,191(2014).
120. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C72**, 2039 (2012).
121. F. Abe *et al.* (CDF Collab.), Phys. Rev. **D50**, 2966 (1994).
122. A. Abulencia *et al.* (CDF Collab.), Phys. Rev. **D73**, 032003 (2006).
123. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C75**, 75 (2015).
124. CMS Collab., CMS PAS TOP-14-001.
125. V.M. Abazov *et al.* (DØ Collab.), Nature **429**, 638 (2004).
126. K. Kondo *et al.*, J. Phys. Soc. Jpn. **G62**, 1177 (1993).
127. R.H. Dalitz and G.R. Goldstein, Phys. Rev. **D45**, 1531 (1992); Phys. Lett. **B287**, 225 (1992); Proc. Royal Soc. London **A445**, 2803 (1999).
128. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **113**, 032002 (2014).
129. L. Sonnenschein, Phys. Rev. **D73**, 054015 (2006).
130. CMS Collab., CMS-PAS-TOP-14-014.
131. CMS Collab., CMS PAS TOP-14-010.
132. B. Abbot *et al.* (DØ Collab.), Phys. Rev. **D60**, 052001 (1999); F. Abe *et al.* (CDF Collab.), Phys. Rev. Lett. **82**, 271 (1999).
133. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D92**, 032003 (2015).
134. CMS Collab., [arXiv:1509.04044](#).
135. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C75**, 158 (2015).
136. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D90**, 091101R, (2014).
137. T. Aaltonen *et al.* (CDF Collab.), Phys. Lett. **B698**, 371 (2011).
138. CMS Collab., CMS PAS TOP-12-030.
139. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D80**, 051104, (2009).
140. V.M. Abazov *et al.* (DØ Collab.) Phys. Rev. Lett. **100**, 192004, (2008); S. Chatrchyan, (CMS Collab.), Phys. Lett. **B728**, 496, (2013); V.M. Abazov *et al.* (DØ Collab.) Phys. Lett. **B703**, 422, (2011); ATLAS Collab., ATLAS-CONF-2011-054; ; U. Langenfeld, S. Moch, and P. Uwer, Phys. Rev. **D80**, 054009 (2009).
141. G. Aad *et al.* (ATLAS Collab.) J. High Energy Phys.1510,121(2015).
142. CMS Collab., CMS PAS TOP-014-015.
143. CDF Collab., CDF conference note 11080 (2014).
144. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D91**, 112003 (2015).
145. The Tevatron Electroweak Working Group, For the CDF and DØ Collab., [arXiv:1407.2682](#).
146. ATLAS & CMS Collabs., ATLAS-CONF-2013-102, CMS PAS TOP-13-005.
147. A.H. Hoang and J.W. Stewart, Nucl. Phys. (Proc. Supp.) **185**, 220 (2008).
148. G. Aad *et al.* (ATLAS Collab.), Phys. Lett. **B716**, 1 (2012).
149. S. Chatrchyan *et al.* (CMS Collab.), Phys. Lett. **B716**, 30 (2012).
150. G. Degrossi, *et al.* J. High Energy Phys.08,98,(2012).
151. S. Alekhin, A. Djouadi, and S. Moch., Phys. Lett. **B716**, 214, (2012).
152. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D87**, 052013 (2013).
153. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D84**, 052005 (2011).
154. G. Aad *et al.* (ATLAS Collab.), Phys. Lett. **B728**, 363 (2014).
155. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.06,109(2012).
156. CMS Collab., CMS-PAS-TOP-12-031.
157. G. Mahlon and S. Parke, Phys. Rev. **D53**, 4886 (1996); G. Mahlon and S. Parke, Phys. Lett. **B411**, 173 (1997).
158. G.R. Goldstein, in *Spin 96: Proceedings of the 12th International Symposium on High Energy Spin Physics*, Amsterdam, 1996, ed. C.W. Jager (World Scientific, Singapore, 1997), p. 328.
159. T. Stelzer and S. Willenbrock, Phys. Lett. **B374**, 169 (1996).
160. W. Bernreuther *et al.* Nucl. Phys. **B690**, 81 (2004).
161. A. Brandenburg, Z.G. Si, & P. Uwer, Phys. Lett. **B539**, 235 (2002).
162. CDF Collab., CDF conference note 10719 (2011).
163. CDF Collab., CDF conference note 10211 (2010).
164. V.M. Abazov *et al.* (DØ Collab.) Phys. Rev. Lett. **108**, 032004, (2012).
165. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **107**, 032001 (2011).
166. V.M. Abazov *et al.* (DØ Collab.), Phys. Lett. **B702**, 16 (2011).
167. G. Mahlon & S.J. Parke, Phys. Rev. **D81**, 074024,2010.
168. G. Aad *et al.* (ATLAS Collab.) Phys. Rev. **D90**, 112016 (2014).
169. W. Bernreuther & Z.G. Si, Nucl. Phys. **B837**, 90 (2010).
170. G. Aad *et al.* (ATLAS Collab.) Phys. Rev. Lett. **114**, 142001 (2015).
171. S. Chatrchyan *et al.* (CMS Collab.) Phys. Rev. Lett. **112**, 182001 (2014).
172. CMS Collab., CMS PAS TOP-13-015.
173. CMS Collab., CMS-PAS-TOP-14-023.
174. G. Aad *et al.* (ATLAS Collab.) Phys. Rev. Lett. **111**, 232002 (2013).
175. A. Falk and M. Peskin, Phys. Rev. **D49**, 3320 (1994).

- 
176. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **111**, 202001 (2013).
177. V.M. Abazov *et al.* (DØ Collab.) Phys. Rev. **D85**, 091104 (2012).
178. V. Khachatryan *et al.* (CMS Collab.) Phys. Lett. **B736**, 33 (2014).
179. G.L. Kane, G.A. Ladinsky, and C.P. Yuan, Phys. Rev. **D45**, 124 (1992).
180. A. Czarnecki, J.G. Korner, and J.H. Piclum, Phys. Rev. **D81**, 111503 (2010).
181. CDF and DØ Collab., Phys. Rev. **D85**, 071106 (2012).
182. CDF Collab., Phys. Rev. **D87**, 031103 (2013).
183. ATLAS Collab., J. High Energy Phys.1206,088(2012).
184. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.10,167(2013).
185. ATLAS and CMS Collab., ATLAS-CONF-2013-033, CMS-PAS-TOP-11-025.
186. CMS Collab., CMS-PAS-TOP-13-008.
187. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.01,053(2015).
188. CMS Collab., CMS-PAS-TOP-14-017.
189. D. Choudhury, T.M.P. Tait, and C.E.M. Wagner, Phys. Rev. **D65**, 053002 (2002).
190. D. Chang, W.F. Chang, and E. Ma, Phys. Rev. **D59**, 091503 (1999), Phys. Rev. **D61**, 037301 (2000).
191. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **98**, 041801 (2007).
192. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D88**, 032003 (2013).
193. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. **D90**, 051101 (2014).
194. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.11,031(2011).
195. CMS Collab., CMS-PAS-TOP-11-031.
196. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D84**, 031104 (2011).
197. ATLAS Collab., ATLAS-CONF-2011-153.
198. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D91**, 072007 (2015).
199. CMS Collab., CMS-PAS-TOP-13-011.
200. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **110**, 172002 (2013).
201. ATLAS Collab., ATLAS-CONF-2012-126.
202. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.1511,172(2015).
203. CMS Collab., [arXiv:1510.01131](#).
204. S.P. Martin, [hep-ph/9709356](#) (1997).
205. C.T. Hill, Phys. Lett. **B266**, 419 (1991).
206. C.T. Hill, Phys. Lett. **B345**, 483 (1995).
207. V. Khachatryan, *et al.* (CMS Collab.), J. High Energy Phys.11,154(2014).
208. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.08,105(2015).
209. T. Plehn and T.M.P. Tait, J. Phys. **G36**, 075001 (2009).
210. C. Zhang and S. Willenbrock, Phys. Rev. **D83**, 034006 (2011).
211. J.A. Aguilar-Saavedra, Nucl. Phys. **B843**, 638 (2011).
212. G. Durieux, F. Maltoni, and C. Zhang, Phys. Rev. **D91**, 7 (2015).
213. V. Barger, T. Han, and D.G.E. Walker, Phys. Rev. Lett. **100**, 031801 (2008).
214. D. Choudhury *et al.*, Phys. Lett. **B657**, 69 (2007).
215. R. Frederix and F. Maltoni, J. High Energy Phys.01,047(2009).
216. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **110**, 121802 (2013).
217. D. Acosta *et al.* (DØ Collab.), Phys. Rev. Lett. **94**, 211801 (2005).
218. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D84**, 072003 (2011).
219. A. Altheimer *et al.*, J. Phys. **G39**, 063001 (2012).
220. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. **D87**, 072002 (2013).
221. S. Chatrchyan *et al.* (CMS Collab.), J. High Energy Phys.1212,015(2012).
222. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **111**, 211804 (2013), Phys. Rev. Lett. **112**, 119903 (2014).
223. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. **D93**, 012001 (2016).
224. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D88**, 012004 (2013).
225. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.1508,148(2015).
226. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.1301,116(2013).
227. T. Plehn *et al.*, J. High Energy Phys.1010,078(2010).
228. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D86**, 091103 (2012).
229. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.11,104(2014).
230. CMS Collab., CMS-PAS-TOP-12-042 (2013).
231. V. Khachatryan *et al.* (CMS Collab.), J. High Energy Phys.06,121(2015).
232. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. Lett. **108**, 041805 (2012).
233. CMS Collab., CMS-PAS-TOP-14-007.
234. CMS Collab., CMS-PAS-TOP-14-003.
235. G. Aad *et al.* (ATLAS Collab.), [arxiv:1509.00294](#).
236. G. Aad *et al.* (ATLAS Collab.), Phys. Lett. **B712**, 351 (2012).
237. A. Heister *et al.* (ALEPH Collab.), Phys. Lett. **B543**, 173 (2002); J. Abdallah *et al.* (DELPHI Collab.), Phys. Lett. **B590**, 21 (2004); P. Achard *et al.* (L3 Collab.), Phys. Lett. **B549**, 290 (2002); G. Abbiendi *et al.* (OPAL Collab.), Phys. Lett. **B521**, 181 (2001).
238. F.D. Aaron *et al.* (H1 Collab.), Phys. Lett. **B678**, 450 (2009).
239. H. Abramowics *et al.* (ZEUS Collab.), Phys. Lett. **B708**, 27 (2012).
240. M. Beneke *et al.*, [hep-ph/0003033](#), in *Proceedings of 1999 CERN Workshop on Standard Model Physics (and more) at the LHC*, G. Altarelli and M.L. Mangano eds.
241. V.F. Obraztsov, S.R. Slabospitsky, and O.P. Yushchenko, Phys. Lett. **B426**, 393 (1998).
242. T. Carli, D. Dannheim, and L. Bellagamba, Mod. Phys. Lett. **A19**, 1881 (2004).

# Quark Particle Listings

*t*

243. V.M. Abazov *et al.* (DØ Collab.), Phys. Rev. Lett. **102**, 092002 (2009).
244. V.M. Abazov *et al.* (DØ Collab.), DØ conference note 5838 (2009).
245. V.M. Abazov *et al.* (DØ Collab.), Phys. Lett. **B708**, 21 (2012).
246. ATLAS Collab., ATLAS-CONF-2013-032.
247. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **101**, 192002 (2009).
248. V.M. Abazov *et al.* (DØ Collab.), Phys. Lett. **B701**, 313 (2011).
249. V.M. Abazov *et al.* (DØ Collab.), Phys. Lett. **B693**, 81 (2010).
250. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **112**, 171802 (2014).
251. G. Aad *et al.* (ATLAS Collab.), Eur. Phys. J. **C76**, 12 (2016).
252. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.1209,139(2012).
253. CMS Collab., CMS-PAS-TOP-13-017.
254. CMS Collab., CMS-PAS-TOP-14-019.
255. G. Aad *et al.* (ATLAS Collab.), J. High Energy Phys.1512,061(2015).

## *t*-QUARK MASS

We first list the direct measurements of the top quark mass which employ the event kinematics and then list the measurements which extract a top quark mass from the measured  $t\bar{t}$  cross-section using theory calculations. A discussion of the definition of the top quark mass in these measurements can be found in the review "The Top Quark."

OUR EVALUATION of  $173.21 \pm 0.51 \pm 0.71$  GeV is an average of published top mass measurements from Tevatron Runs. The first combination of the top-quark mass measurements, including some unpublished data, has been performed by the CDF and DØ experiments at the Tevatron and ATLAS and CMS experiments at the LHC. The resulting combined top-quark mass is  $173.34 \pm 0.27 \pm 0.71$  GeV, consistent with Tevatron average. The latest Tevatron average,  $174.34 \pm 0.37 \pm 0.52$  GeV, was provided by the Tevatron Electroweak Working Group (TEVEWWG). It takes correlated uncertainties into account and has a  $\chi^2$  of 10.8 for 11 degrees of freedom.

For earlier search limits see PDG 96, Physical Review **D54** 1 (1996). We no longer include a compilation of indirect top mass determinations from Standard Model Electroweak fits in the Listings (our last compilation can be found in the Listings of the 2007 partial update). For a discussion of current results see the reviews "The Top Quark" and "Electroweak Model and Constraints on New Physics."

## *t*-Quark Mass (Direct Measurements)

The following measurements extract a *t*-quark mass from the kinematics of  $t\bar{t}$  events. They are sensitive to the top quark mass used in the MC generator that is usually interpreted as the pole mass, but the theoretical uncertainty in this interpretation is hard to quantify. See the review "The Top Quark" and references therein for more information.

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b><math>173.21 \pm 0.51 \pm 0.71</math></b>	<b>OUR EVALUATION</b>		See comments in the header above.
$173.32 \pm 1.36 \pm 0.85$	<sup>1</sup> ABAZOV	16 D0	$\ell\ell + \cancel{E}_T + \geq 2j$ ( $\geq 2b$ )
$175.1 \pm 1.4 \pm 1.2$	<sup>2</sup> AAD	15AW ATLS	small $\cancel{E}_T$ , $\geq 6$ jets (2 <i>b</i> -tag)
$172.99 \pm 0.48 \pm 0.78$	<sup>3</sup> AAD	15BF ATLS	$\ell + \text{jets}$ and dilepton
$171.5 \pm 1.9 \pm 2.5$	<sup>4</sup> AALTONEN	15D CDF	$\ell\ell + \cancel{E}_T + \geq 2j$
$175.07 \pm 1.19 \pm 1.55$	<sup>5</sup> AALTONEN	14N CDF	small $\cancel{E}_T$ , 6-8 jets ( $\geq 1b$ -tag)
$174.98 \pm 0.58 \pm 0.49$	<sup>6</sup> ABAZOV	14C D0	$\ell + \cancel{E}_T + 4$ jets ( $\geq 1b$ -tag)
$173.49 \pm 0.69 \pm 1.21$	<sup>7</sup> CHATRCHYAN	14C CMS	$\geq 6$ jets ( $\geq 2b$ -tag)
$173.93 \pm 1.64 \pm 0.87$	<sup>8</sup> AALTONEN	13H CDF	$\cancel{E}_T + \geq 4$ jets ( $\geq 1b$ )
$173.9 \pm 0.9 \pm 1.7$	<sup>9</sup> CHATRCHYAN	13S CMS	$\ell\ell + \cancel{E}_T + \geq 2b$ -tag (MT2( $T$ ))
$172.85 \pm 0.71 \pm 0.85$	<sup>10</sup> AALTONEN	12AI CDF	$\ell + \cancel{E}_T + \geq 4j$ (0,1,2 <i>b</i> ) template
$172.7 \pm 9.3 \pm 3.7$	<sup>11</sup> AALTONEN	12AL CDF	$\tau_h + \cancel{E}_T + 4j$ ( $\geq 1b$ )
$173.9 \pm 1.9 \pm 1.6$	<sup>12</sup> ABAZOV	12AB D0	$\ell\ell + \cancel{E}_T + \geq 2j$ ( $\nu$ WT+MWT)

$172.5 \pm 0.4 \pm 1.5$	<sup>13</sup> CHATRCHYAN	12BA CMS	$\ell\ell + \cancel{E}_T + \geq 2j$ ( $\geq 1b$ ), AMWT
$173.49 \pm 0.43 \pm 0.98$	<sup>14</sup> CHATRCHYAN	12BP CMS	$\ell + \cancel{E}_T + \geq 4j$ ( $\geq 2b$ )
$173.0 \pm 1.2$	<sup>15</sup> AALTONEN	10AE CDF	$\ell + \cancel{E}_T + 4$ jets ( $\geq 1b$ -tag), ME method
$170.7 \pm 6.3 \pm 2.6$	<sup>16</sup> AALTONEN	10D CDF	$\ell + \cancel{E}_T + 4$ jets ( <i>b</i> -tag)
$180.1 \pm 3.6 \pm 3.9$	<sup>17,18</sup> ABAZOV	04G D0	lepton + jets
$176.1 \pm 5.1 \pm 5.3$	<sup>19</sup> AFFOLDER	01 CDF	lepton + jets
$167.4 \pm 10.3 \pm 4.8$	<sup>20,21</sup> ABE	99B CDF	dilepton
$168.4 \pm 12.3 \pm 3.6$	<sup>18</sup> ABBOTT	98D D0	dilepton
$186 \pm 10 \pm 5.7$	<sup>20,22</sup> ABE	97R CDF	6 or more jets
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$174.5 \pm 0.6 \pm 2.3$	<sup>23</sup> AAD	12I ATLS	$\ell + \cancel{E}_T + \geq 4$ jets ( $\geq 1b$ ), MT
$173.18 \pm 0.56 \pm 0.75$	<sup>24</sup> AALTONEN	12AP TEVA	CDF, D0 combination
$172.5 \pm 1.4 \pm 1.5$	<sup>25</sup> AALTONEN	12G CDF	6-8 jets with $\geq 1b$
$173.7 \pm 2.8 \pm 1.5$	<sup>26</sup> ABAZOV	12AB D0	$\ell\ell + \cancel{E}_T + \geq 2j$ ( $\nu$ WT)
$172.4 \pm 1.4 \pm 1.3$	<sup>27</sup> AALTONEN	11AC CDF	$\ell + \cancel{E}_T + 4$ jets ( $\geq 1b$ -tag)
$172.3 \pm 2.4 \pm 1.0$	<sup>28</sup> AALTONEN	11AK CDF	Repl. by AALTONEN 13H
$172.1 \pm 1.1 \pm 0.9$	<sup>29</sup> AALTONEN	11E CDF	$\ell + \text{jets}$ and dilepton
$176.9 \pm 8.0 \pm 2.7$	<sup>30</sup> AALTONEN	11T CDF	$\ell + \cancel{E}_T + 4$ jets ( $\geq 1b$ -tag), $p_T(\ell)$ shape
$174.94 \pm 0.83 \pm 1.24$	<sup>31</sup> ABAZOV	11P D0	$\ell + \cancel{E}_T + 4$ jets ( $\geq 1b$ -tag)
$174.0 \pm 1.8 \pm 2.4$	<sup>32</sup> ABAZOV	11R D0	dilepton + $\cancel{E}_T + \geq 2$ jets
$175.5 \pm 4.6 \pm 4.6$	<sup>33</sup> CHATRCHYAN	11F CMS	dilepton + $\cancel{E}_T + \text{jets}$
$169.3 \pm 2.7 \pm 3.2$	<sup>34</sup> AALTONEN	10C CDF	dilepton + <i>b</i> -tag (MT2+NWA)
$174.8 \pm 2.4 \pm 1.2$	<sup>35</sup> AALTONEN	10E CDF	$\geq 6$ jets, vtx <i>b</i> -tag
$180.5 \pm 12.0 \pm 3.6$	<sup>36</sup> AALTONEN	09AK CDF	$\ell + \cancel{E}_T + \text{jets}$ (soft $\mu$ <i>b</i> -tag)
$172.7 \pm 1.8 \pm 1.2$	<sup>37</sup> AALTONEN	09J CDF	$\ell + \cancel{E}_T + 4$ jets ( <i>b</i> -tag)
$171.1 \pm 3.7 \pm 2.1$	<sup>38</sup> AALTONEN	09K CDF	6 jets, vtx <i>b</i> -tag
$171.9 \pm 1.7 \pm 1.1$	<sup>39</sup> AALTONEN	09L CDF	$\ell + \text{jets}$ , $\ell\ell + \text{jets}$
$171.2 \pm 2.7 \pm 2.9$	<sup>40</sup> AALTONEN	09O CDF	dilepton
$165.5 \pm 3.4 \pm 3.3$	<sup>41</sup> AALTONEN	09X CDF	$\ell\ell + \cancel{E}_T$ ( $\nu\phi$ weighting)
$174.7 \pm 4.4 \pm 2.0$	<sup>42</sup> ABAZOV	09AH D0	dilepton + <i>b</i> -tag ( $\nu$ WT+MWT)
$170.7 \pm 4.2 \pm 3.9$	<sup>43,44</sup> AALTONEN	08C CDF	dilepton, $\sigma_{t\bar{t}}$ constrained
$171.5 \pm 1.8 \pm 1.1$	<sup>45</sup> ABAZOV	08AH D0	$\ell + \cancel{E}_T + 4$ jets
$177.1 \pm 4.9 \pm 4.7$	<sup>46,47</sup> AALTONEN	07 CDF	6 jets with $\geq 1b$ vtx
$172.3 \pm 10.8 \pm 9.6$	<sup>48</sup> AALTONEN	07B CDF	$\geq 4$ jets ( <i>b</i> -tag)
$174.0 \pm 2.2 \pm 4.8$	<sup>49</sup> AALTONEN	07D CDF	$\geq 6$ jets, vtx <i>b</i> -tag
$170.8 \pm 2.2 \pm 1.4$	<sup>50,51</sup> AALTONEN	07I CDF	lepton + jets ( <i>b</i> -tag)
$173.7 \pm 4.4 \pm 2.1$	<sup>47,52</sup> ABAZOV	07F D0	lepton + jets
$176.2 \pm 9.2 \pm 3.9$	<sup>53</sup> ABAZOV	07W D0	dilepton (MWT)
$179.5 \pm 7.4 \pm 5.6$	<sup>53</sup> ABAZOV	07W D0	dilepton ( $\nu$ WT)
$164.5 \pm 3.9 \pm 3.9$	<sup>51,54</sup> ABULENCIA	07D CDF	dilepton
$180.7 \pm 15.5 \pm 13.4$	<sup>55</sup> ABULENCIA	07J CDF	lepton + jets
$170.3 \pm 4.1 \pm 4.5$	<sup>51,56</sup> ABAZOV	06U D0	lepton + jets ( <i>b</i> -tag)
$173.2 \pm 2.6 \pm 2.4$	<sup>57,58</sup> ABULENCIA	06D CDF	lepton + jets
$173.5 \pm 3.7 \pm 3.6$	<sup>44,57</sup> ABULENCIA	06D CDF	lepton + jets
$165.2 \pm 6.1 \pm 3.4$	<sup>51,59</sup> ABULENCIA	06G CDF	dilepton
$170.1 \pm 6.0 \pm 4.1$	<sup>44,60</sup> ABULENCIA	06V CDF	dilepton
$178.5 \pm 13.7 \pm 7.7$	<sup>61,62</sup> ABAZOV	05 D0	6 or more jets
$176.1 \pm 6.6$	<sup>63</sup> AFFOLDER	01 CDF	dilepton, lepton+jets, all-jets
$172.1 \pm 5.2 \pm 4.9$	<sup>64</sup> ABBOTT	99G D0	di-lepton, lepton+jets
$176.0 \pm 6.5$	<sup>21,65</sup> ABE	99B CDF	dilepton, lepton+jets, all-jets
$173.3 \pm 5.6 \pm 5.5$	<sup>18,66</sup> ABBOTT	98R D0	lepton + jets
$175.9 \pm 4.8 \pm 5.3$	<sup>20,67</sup> ABE	98E CDF	lepton + jets
$161 \pm 17 \pm 10$	<sup>20</sup> ABE	98F CDF	dilepton
$172.1 \pm 5.2 \pm 4.9$	<sup>68</sup> BHAT	98B RVUE	dilepton and lepton+jets
$173.8 \pm 5.0$	<sup>69</sup> BHAT	98B RVUE	dilepton, lepton+jets, all-jets
$173.3 \pm 5.6 \pm 6.2$	<sup>18</sup> ABACHI	97E D0	lepton + jets
$199 \pm 19 \pm 21$	ABACHI	95 D0	lepton + jets
$176 \pm 8 \pm 10$	ABE	95F CDF	lepton + <i>b</i> -jet
$174 \pm 10 \pm 13$	ABE	94E CDF	lepton + <i>b</i> -jet

<sup>1</sup> ABAZOV 16 based on  $9.7 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . Employs improved fit to minimize statistical errors and improved jet energy calibration, using lepton + jets mode, which reduces error of jet energy scale. Based on previous determination in ABAZOV 12AB with increased integrated luminosity and improved fit and calibrations.

<sup>2</sup> AAD 15AW based on  $4.6 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . Uses template fits to the ratio of the masses of three-jets (from *t* candidate) and dijets (from *W* candidate). Large background from multijet production is modeled with data-driven methods.

<sup>3</sup> AAD 15BF based on  $4.6 \text{ fb}^{-1}$  in  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$ . Using a three-dimensional template likelihood technique the lepton plus jets ( $\geq 1b$ -tagged) channel gives  $172.33 \pm 0.75 \pm 1.02 \text{ GeV}$ , while exploiting a one dimensional template method using  $m_{\ell\ell}$  the dilepton channel (1 or 2*b*-tags) gives  $173.79 \pm 0.54 \pm 1.30 \text{ GeV}$ . The results are combined.

<sup>4</sup> AALTONEN 15D based on  $9.1 \text{ fb}^{-1}$  of  $p\bar{p}$  data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Uses a template technique to fit a distribution of a variable defined by a linear combination of variables sensitive and insensitive to jet energy scale to optimize reduction of systematic errors. *b*-tagged and non-*b*-tagged events are separately analyzed and combined.

<sup>5</sup> Based on  $9.3 \text{ fb}^{-1}$  of  $p\bar{p}$  data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Multivariate algorithm is used to discriminate signal from backgrounds, and templates are used to measure  $m_t$ .

<sup>6</sup> Based on  $9.7 \text{ fb}^{-1}$  of  $p\bar{p}$  data at  $\sqrt{s} = 1.96 \text{ TeV}$ . A matrix element method is used to calculate the probability of an event to be signal or background, and the overall jet energy scale is constrained *in situ* by  $m_{W\ell}$ . See ABAZOV 15G for further details.

- <sup>7</sup> Based on  $3.54 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . The mass is reconstructed for each event employing a kinematic fit of the jets to a  $t\bar{t}$  hypothesis. The combination with the previous CMS measurements in the dilepton and the lepton+jets channels gives  $173.54 \pm 0.33 \pm 0.96 \text{ GeV}$ .
- <sup>8</sup> Based on  $8.7 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . Events with an identified charged lepton or small  $E_T$  are rejected from the event sample, so that the measurement is statistically independent from those in the  $\ell + \text{jets}$  and all hadronic channels while being sensitive to those events with a  $\tau$  lepton in the final state.
- <sup>9</sup> Based on  $5.0 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . CHATRCHYAN 13s studied events with di-lepton +  $E_T + \geq 2 \text{ b-jets}$ , and looked for kinematical endpoints of MT2, MT2T, and subsystem variables.
- <sup>10</sup> Based on  $8.7 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ . The JES is calibrated by using the dijet mass from the  $W$  boson decay.
- <sup>11</sup> Use the ME method based on  $2.2 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ .
- <sup>12</sup> Combination with the result in  $1 \text{ fb}^{-1}$  of preceding data reported in ABAZOV 09AH as well as the MWT result of ABAZOV 11R with a statistical correlation of 60%.
- <sup>13</sup> Based on  $5.0 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . Uses an analytical matrix weighting technique (AMWT) and full kinematic analysis (KIN).
- <sup>14</sup> Based on  $5.0 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . The first error is statistical and JES combined, and the second is systematic. Ideogram method is used to obtain 2D likelihood for the kinematical fit with two parameters  $m_{\text{top}}$  and JES.
- <sup>15</sup> Based on  $5.6 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . The likelihood calculated using a matrix element method gives  $m_t = 173.0 \pm 0.7(\text{stat}) \pm 0.6(\text{JES}) \pm 0.9(\text{syst}) \text{ GeV}$ , for a total uncertainty of  $1.2 \text{ GeV}$ .
- <sup>16</sup> Based on  $1.9 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . The result is from the measurement using the transverse decay length of  $b$ -hadrons and that using the transverse momentum of the  $W$  decay muons, which are both insensitive to the JES (jet energy scale) uncertainty. OUR EVALUATION uses only the measurement exploiting the decay length significance which yields  $166.9^{+9.5}_{-8.5}(\text{stat}) \pm 2.9(\text{syst}) \text{ GeV}$ . The measurement that uses the lepton transverse momentum is excluded from the average because of a statistical correlation with other samples.
- <sup>17</sup> Obtained by re-analysis of the lepton + jets candidate events that led to ABBOTT 98f. It is based upon the maximum likelihood method which makes use of the leading order matrix elements.
- <sup>18</sup> Based on  $125 \pm 7 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.8 \text{ TeV}$ .
- <sup>19</sup> Based on  $\sim 106 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.8 \text{ TeV}$ .
- <sup>20</sup> Based on  $109 \pm 7 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.8 \text{ TeV}$ .
- <sup>21</sup> See AFFOLDER 01 for details of systematic error re-evaluation.
- <sup>22</sup> Based on the first observation of all hadronic decays of  $t\bar{t}$  pairs. Single  $b$ -quark tagging with jet-shape variable constraints was used to select signal enriched multi-jet events. The updated systematic error is listed. See AFFOLDER 01, appendix C.
- <sup>23</sup> AAD 12i based on  $1.04 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . Uses 2d-template analysis (MT) with  $m_t$  and jet energy scale factor (JSF) from  $m_W$  mass fit.
- <sup>24</sup> Combination based on up to  $5.8 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ .
- <sup>25</sup> Based on  $5.8 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ . The quoted systematic error is the sum of JES( $\pm 1.0$ ) and systematic( $\pm 1.1$ ) uncertainties. The measurement is performed with a likelihood fit technique which simultaneously determines  $m_t$  and JES.
- <sup>26</sup> Based on  $4.3 \text{ fb}^{-1}$  of data in  $p$ - $\bar{p}$  collisions at  $1.96 \text{ TeV}$ . The measurement reduces the JES uncertainty by using the single lepton channel study of ABAZOV 11P.
- <sup>27</sup> Based on  $3.2 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is from statistics and JES combined, and the latter is from the other systematic uncertainties. The result is obtained using an unbinned maximum likelihood method where the top quark mass and the JES are measured simultaneously, with  $\Delta_{JES} = 0.3 \pm 0.3(\text{stat})$ .
- <sup>28</sup> Based on  $5.7 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . Events with an identified charged lepton or small  $E_T$  are rejected from the event sample, so that the measurement is statistically independent from those in the  $\ell + \text{jets}$  and all hadronic channels while being sensitive to those events with a  $\tau$  lepton in the final state. Supersedes AALTONEN 07B.
- <sup>29</sup> AALTONEN 11E based on  $5.6 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . Employs a multi-dimensional template likelihood technique where the lepton plus jets (one or two  $b$ -tags) channel gives  $172.2 \pm 1.2 \pm 0.9 \text{ GeV}$  while the dilepton channel yields  $170.3 \pm 2.0 \pm 3.1 \text{ GeV}$ . The results are combined. OUR EVALUATION includes the measurement in the dilepton channel only.
- <sup>30</sup> Uses a likelihood fit of the lepton  $p_T$  distribution based on  $2.7 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>31</sup> Based on  $3.6 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . ABAZOV 11P reports  $174.94 \pm 0.83 \pm 0.78 \pm 0.96 \text{ GeV}$ , where the first uncertainty is from statistics, the second from JES, and the last from other systematic uncertainties. We combine the JES and systematic uncertainties. A matrix-element method is used where the JES uncertainty is constrained by the  $W$  mass. ABAZOV 11P describes a measurement based on  $2.6 \text{ fb}^{-1}$  that is combined with ABAZOV 08AH, which employs an independent  $1 \text{ fb}^{-1}$  of data.
- <sup>32</sup> Based on a matrix-element method which employs  $5.4 \text{ fb}^{-1}$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . Superseded by ABAZOV 12AB.
- <sup>33</sup> Based on  $36 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$ . A Kinematic Method using  $b$ -tagging and an analytical Matrix Weighting Technique give consistent results and are combined. Superseded by CHATRCHYAN 12BA.
- <sup>34</sup> Based on  $3.4 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . The result is obtained by combining the MT2 variable method and the NVA (Neutrino Weighting Algorithm). The MT2 method alone gives  $m_t = 168.0^{+4.8}_{-4.0}(\text{stat}) \pm 2.9(\text{syst}) \text{ GeV}$  with smaller systematic error due to small JES uncertainty.
- <sup>35</sup> Based on  $2.9 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is from statistics and JES uncertainty, and the latter is from the other systematics. Neural-network-based kinematical selection of 6 highest  $E_T$  jets with a vtx  $b$ -tag is used to distinguish signal from background. Superseded by AALTONEN 12G.
- <sup>36</sup> Based on  $2 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The top mass is obtained from the measurement of the invariant mass of the lepton ( $e$  or  $\mu$ ) from  $W$  decays and the soft  $\mu$  in  $b$ -jet. The result is insensitive to jet energy scaling.
- <sup>37</sup> Based on  $1.9 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is from statistics and jet energy scale uncertainty, and the latter is from the other systematics. Matrix element method with effective propagators.
- <sup>38</sup> Based on  $943 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is from statistical and jet-energy-scale uncertainties, and the latter is from other systematics. AALTONEN 09k selected 6 jet events with one or more vertex  $b$ -tags and used the tree-level matrix element to construct template models of signal and background.

- <sup>39</sup> Based on  $1.9 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is from statistical and jet-energy-scale (JES) uncertainties, and the second is from other systematics. Events with lepton + jets and those with dilepton + jets were simultaneously fit to constrain  $m_t$  and JES. Lepton + jets data only give  $m_t = 171.8 \pm 2.2 \text{ GeV}$ , and dilepton data only give  $m_t = 171.2^{+5.3}_{-5.1} \text{ GeV}$ .
- <sup>40</sup> Based on  $2 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Matrix Element method. Optimal selection criteria for candidate events with two high  $p_T$  leptons, high  $E_T$ , and two or more jets with and without  $b$ -tag are obtained by neural network with neuroevolution technique to minimize the statistical error of  $m_t$ .
- <sup>41</sup> Based on  $2.9 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Mass  $m_t$  is estimated from the likelihood for the eight-fold kinematical solutions in the plane of the azimuthal angles of the two neutrino momenta.
- <sup>42</sup> Based on  $1 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Events with two identified leptons, and those with one lepton plus one isolated track and a  $b$ -tag were used to constrain  $m_t$ . The result is a combination of the  $\nu$ WT ( $\nu$  Weighting Technique) result of  $176.2 \pm 4.8 \pm 2.1 \text{ GeV}$  and the MWT (Matrix-element Weighting Technique) result of  $173.2 \pm 4.9 \pm 2.0 \text{ GeV}$ .
- <sup>43</sup> Reports measurement of  $170.7^{+4.2}_{-3.9} \pm 2.6 \pm 2.4 \text{ GeV}$  based on  $1.2 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The last error is due to the theoretical uncertainty on  $\sigma_{t\bar{t}}$ . Without the cross-section constraint a top mass of  $169.7^{+5.2}_{-4.9} \pm 3.1 \text{ GeV}$  is obtained.
- <sup>44</sup> Template method.
- <sup>45</sup> Result is based on  $1 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is from statistics and jet energy scale uncertainty, and the latter is from the other systematics.
- <sup>46</sup> Based on  $310 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>47</sup> Ideogram method.
- <sup>48</sup> Based on  $311 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Events with 4 or more jets with  $E_T > 15 \text{ GeV}$ , significant missing  $E_T$ , and secondary vertex  $b$ -tag are used in the fit. About 44% of the signal acceptance is from  $\tau\nu + 4 \text{ jets}$ . Events with identified  $e$  or  $\mu$  are vetoed to provide a statistically independent measurement.
- <sup>49</sup> Based on  $1.02 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Superseded by AALTONEN 12G.
- <sup>50</sup> Based on  $955 \text{ pb}^{-1}$  of data  $\sqrt{s} = 1.96 \text{ TeV}$ .  $m_t$  and JES (Jet Energy Scale) are fitted simultaneously, and the first error contains the JES contribution of  $1.5 \text{ GeV}$ .
- <sup>51</sup> Matrix element method.
- <sup>52</sup> Based on  $425 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error is a combination of statistics and JES (Jet Energy Scale) uncertainty, which has been measured simultaneously to give  $JES = 0.989 \pm 0.029(\text{stat})$ .
- <sup>53</sup> Based on  $370 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Combined result of MWT (Matrix-element Weighting Technique) and  $\nu$ WT ( $\nu$  Weighting Technique) analyses is  $178.1 \pm 6.7 \pm 4.8 \text{ GeV}$ .
- <sup>54</sup> Based on  $1.0 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . ABULENCIA 07D improves the matrix element description by including the effects of initial-state radiation.
- <sup>55</sup> Based on  $695 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The transverse decay length of the  $b$  hadron is used to determine  $m_t$ , and the result is free from the JES (jet energy scale) uncertainty.
- <sup>56</sup> Based on  $\sim 400 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . The first error includes statistical and systematic jet energy scale uncertainties, the second error is from the other systematics. The result is obtained with the  $b$ -tagging information. The result without  $b$ -tagging is  $169.2^{+5.0}_{-7.4} \pm 1.5 \pm 1.4 \text{ GeV}$ . Superseded by ABAZOV 08AH.
- <sup>57</sup> Based on  $318 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>58</sup> Dynamical likelihood method.
- <sup>59</sup> Based on  $340 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>60</sup> Based on  $360 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>61</sup> Based on  $110.2 \pm 5.8 \text{ pb}^{-1}$  at  $\sqrt{s} = 1.8 \text{ TeV}$ .
- <sup>62</sup> Based on the all hadronic decays of  $t\bar{t}$  pairs. Single  $b$ -quark tagging via the decay chain  $b \rightarrow c \rightarrow \mu$  was used to select signal enriched multijet events. The result was obtained by the maximum likelihood method after bias correction.
- <sup>63</sup> Obtained by combining the measurements in the lepton + jets [AFFOLDER 01], all-jets [ABE 97R, ABE 99B], and dilepton [ABE 99B] decay topologies.
- <sup>64</sup> Obtained by combining the D0 result  $m_t(\text{GeV}) = 168.4 \pm 12.3 \pm 3.6$  from 6 di-lepton events (see also ABBOTT 98d) and  $m_t(\text{GeV}) = 173.3 \pm 5.6 \pm 5.5$  from lepton+jets events (ABBOTT 98f).
- <sup>65</sup> Obtained by combining the CDF results of  $m_t(\text{GeV}) = 167.4 \pm 10.3 \pm 4.8$  from 8 dilepton events,  $m_t(\text{GeV}) = 175.9 \pm 4.8 \pm 5.3$  from lepton+jets events (ABE 98E), and  $m_t(\text{GeV}) = 186.0 \pm 10.0 \pm 5.7$  from all-jet events (ABE 97R). The systematic errors in the latter two measurements are changed in this paper.
- <sup>66</sup> See ABAZOV 04G.
- <sup>67</sup> The updated systematic error is listed. See AFFOLDER 01, appendix C.
- <sup>68</sup> Obtained by combining the D0 results of  $m_t(\text{GeV}) = 168.4 \pm 12.3 \pm 3.6$  from 6 dilepton events and  $m_t(\text{GeV}) = 173.3 \pm 5.6 \pm 5.5$  from 77 lepton+jets events.
- <sup>69</sup> Obtained by combining the D0 results from dilepton and lepton+jets events, and the CDF results (ABE 99B) from dilepton, lepton+jets events, and all-jet events.

### t-Quark $\overline{MS}$ Mass from Cross-Section Measurements

The top quark  $\overline{MS}$  or pole mass can be extracted from a measurement of  $\sigma(t\bar{t})$  by using theory calculations. We quote below the  $\overline{MS}$  mass. See the review "The Top Quark" and references therein for more information.

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b>160.0<sup>+4.8</sup><sub>-4.3</sub></b>	<sup>1</sup> ABAZOV	11s D0	$\sigma(t\bar{t}) + \text{theory}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	<sup>2</sup> ABAZOV	09AG D0	cross sects, theory + exp
	<sup>3</sup> ABAZOV	09R D0	cross sects, theory + exp

# Quark Particle Listings

*t*

- <sup>1</sup> Based on 5.3 fb<sup>-1</sup> in *p* $\bar{p}$  collisions at  $\sqrt{s}$  = 1.96 TeV. ABAZOV 11s uses the measured *t* $\bar{t}$  production cross section of 8.13<sup>+1.02</sup><sub>-0.90</sub> pb [ABAZOV 11E] in the lepton plus jets channel to obtain the top quark  $\overline{MS}$  mass by using an approximate NNLO computation (MOCH 08, LANGENFELD 09). The corresponding top quark pole mass is 167.5<sup>+5.4</sup><sub>-4.9</sub> GeV. A different theory calculation (AHRENS 10, AHRENS 10A) is also used and yields  $m_t^{\overline{MS}} = 154.5^{+5.0}_{-4.3}$  GeV.
- <sup>2</sup> Based on 1 fb<sup>-1</sup> of data at  $\sqrt{s}$  = 1.96 TeV. Uses the  $\ell$  + jets,  $\ell\ell$ , and  $\ell\tau$  + jets channels. ABAZOV 09AG extract the pole mass of the top quark using two different calculations that yield 169.1<sup>+5.9</sup><sub>-5.2</sub> GeV (MOCH 08, LANGENFELD 09) and 168.2<sup>+5.9</sup><sub>-5.4</sub> GeV (KIDONAKIS 08).
- <sup>3</sup> Based on 1 fb<sup>-1</sup> of data at  $\sqrt{s}$  = 1.96 TeV. Uses the  $\ell\ell$  and  $\ell\tau$  + jets channels. ABAZOV 09R extract the pole mass of the top quark using two different calculations that yield 173.3<sup>+9.8</sup><sub>-8.6</sub> GeV (MOCH 08, LANGENFELD 09) and 171.5<sup>+9.9</sup><sub>-8.8</sub> GeV (CACCIARI 08).

## t-Quark Pole Mass from Cross-Section Measurements

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b>174.2<math>\pm</math>1.4 OUR AVERAGE</b>			
173.7 <sup>+2.3</sup> <sub>-2.1</sub>	<sup>1</sup> AAD	15BWATLS	$\ell$ + $\cancel{E}_T$ + $\geq$ 5j (2b-tag)
172.9 <sup>+2.5</sup> <sub>-2.6</sub>	<sup>2</sup> AAD	14AY ATLS	$pp$ at $\sqrt{s}$ = 7, 8 TeV
176.7 <sup>+3.0</sup> <sub>-2.8</sub>	<sup>3</sup> CHATRCHYAN 14	CMS	$pp$ at $\sqrt{s}$ = 7 TeV

- <sup>1</sup> AAD 15BW based on 4.6 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s}$  = 7 TeV. Uses normalized differential cross section for *t* $\bar{t}$  + 1 jet as a function of the inverse of the invariant mass of the *t* $\bar{t}$  + 1 jet system. The measured cross section is corrected to the parton level. Then a fit to the data using NLO + parton shower prediction is performed.
- <sup>2</sup> Used  $\sigma(t\bar{t})$  for  $e\mu$  events. The result is a combination of the measurements  $m_t = 171.4 \pm 2.6$  GeV based on 4.6 fb<sup>-1</sup> of data at 7 TeV and  $m_t = 174.1 \pm 2.6$  GeV based on 20.3 fb<sup>-1</sup> of data at 8 TeV.
- <sup>3</sup> Used  $\sigma(t\bar{t})$  from  $pp$  collisions at  $\sqrt{s}$  = 7 TeV measured in CHATRCHYAN 12AX to obtain  $m_t(\text{pole})$  for  $\alpha_s(m_Z) = 0.1184 \pm 0.0007$ . The errors have been corrected in KHACHATRYAN 14K.

$$m_t - m_{\bar{t}}$$

Test of *CPT* conservation. OUR AVERAGE assumes that the systematic uncertainties are uncorrelated.

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
<b>-0.2 <math>\pm</math> 0.5 OUR AVERAGE</b>	Error includes scale factor of 1.1.		
0.67 $\pm$ 0.61 $\pm$ 0.41	<sup>1</sup> AAD	14 ATLS	$\ell$ + $\cancel{E}_T$ + $\geq$ 4j ( $\geq$ 2 b-tags)
-1.95 $\pm$ 1.11 $\pm$ 0.59	<sup>2</sup> AALTONEN	13E CDF	$\ell$ + $\cancel{E}_T$ + $\geq$ 4j (0,1,2 b-tags)
-0.44 $\pm$ 0.46 $\pm$ 0.27	<sup>3</sup> CHATRCHYAN 12Y	CMS	$\ell$ + $\cancel{E}_T$ + $\geq$ 4j
0.8 $\pm$ 1.8 $\pm$ 0.5	<sup>4</sup> ABAZOV	11T D0	$\ell$ + $\cancel{E}_T$ + 4 jets ( $\geq$ 1 b-tag)
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-3.3 $\pm$ 1.4 $\pm$ 1.0	<sup>5</sup> AALTONEN	11K CDF	Repl. by AALTONEN 13E
3.8 $\pm$ 3.4 $\pm$ 1.2	<sup>6</sup> ABAZOV	09AA D0	$\ell$ + $\cancel{E}_T$ + 4 jets ( $\geq$ 1 b-tag)
<sup>1</sup> Based on 4.7 fb <sup>-1</sup> of $pp$ data at $\sqrt{s}$ = 7 TeV and an average top mass of 172.5 GeV/ <i>c</i> <sup>2</sup> .			
<sup>2</sup> Based on 8.7 fb <sup>-1</sup> of $p\bar{p}$ collisions at $\sqrt{s}$ = 1.96 TeV and an average top mass of 172.5 GeV/ <i>c</i> <sup>2</sup> .			
<sup>3</sup> Based on 4.96 fb <sup>-1</sup> of $pp$ data at $\sqrt{s}$ = 7 TeV. Based on the fitted $m_t$ for $\ell^+$ and $\ell^-$ events using the Ideogram method.			
<sup>4</sup> Based on a matrix-element method which employs 3.6 fb <sup>-1</sup> in $p\bar{p}$ collisions at $\sqrt{s}$ = 1.96 TeV.			
<sup>5</sup> Based on a template likelihood technique which employs 5.6 fb <sup>-1</sup> in $p\bar{p}$ collisions at $\sqrt{s}$ = 1.96 TeV.			
<sup>6</sup> Based on 1 fb <sup>-1</sup> of data in $p\bar{p}$ collisions at $\sqrt{s}$ = 1.96 TeV.			

## t-quark DECAY WIDTH

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.41 <math>\pm</math> 0.19 <math>\pm</math> 0.15 OUR AVERAGE</b>		Error includes scale factor of 1.4.		
1.36 $\pm$ 0.02 $\pm$ 0.14 $\pm$ 0.11		<sup>1</sup> KHACHATRY...14E	CMS	$\ell\ell$ + $\cancel{E}_T$ +2-4jets (0-2b-tag)
2.00 $\pm$ 0.47 $\pm$ 0.43		<sup>2</sup> ABAZOV	12T D0	$\Gamma(t \rightarrow bW)/\Gamma(t \rightarrow bW)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 6.38	95	<sup>3</sup> AALTONEN	13Z CDF	$\ell$ + $\cancel{E}_T$ + $\geq$ 4j ( $\geq$ 0 b), direct
1.99 $\pm$ 0.69 $\pm$ 0.55		<sup>4</sup> ABAZOV	11B D0	Repl. by ABAZOV 12T
> 1.21	95	<sup>4</sup> ABAZOV	11B D0	$\Gamma(t \rightarrow Wb)$
< 7.6	95	<sup>5</sup> AALTONEN	10AC CDF	$\ell$ + jets, direct
<13.1	95	<sup>6</sup> AALTONEN	09M CDF	$m_t(\text{rec})$ distribution

- <sup>1</sup> Based on 19.7 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s}$  = 8 TeV. The result is obtained by combining the measurement of  $R = \Gamma(t \rightarrow Wb)/\Gamma(t \rightarrow Wq (q=b,s,d))$  and a previous CMS measurement of the *t*-channel single top production cross section of CHATRCHYAN 12Bq, by using the theoretical calculation of  $\Gamma(t \rightarrow Wb)$  for  $m_t = 172.5$  GeV.
- <sup>2</sup> Based on 5.4 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV.  $\Gamma(t \rightarrow bW) = 1.87^{+0.44}_{-0.40}$  GeV is obtained from the observed *t*-channel single top quark production cross section, whereas  $\Gamma(t \rightarrow bW) = 0.90 \pm 0.04$  is used assuming  $\sum_q \Gamma(t \rightarrow qW) = 1$ . The result is valid for  $m_t = 172.5$  GeV. See the paper for the values for  $m_t = 170$  or 175 GeV.
- <sup>3</sup> Based on 8.7 fb<sup>-1</sup> of data. The two sided 68% CL interval is 1.10 GeV <  $\Gamma_t$  < 4.05 GeV for  $m_t = 172.5$  GeV.

- <sup>4</sup> Based on 2.3 fb<sup>-1</sup> in  $p\bar{p}$  collisions at  $\sqrt{s}$  = 1.96 TeV. ABAZOV 11B extracted  $\Gamma_t$  from the partial width  $\Gamma(t \rightarrow Wb) = 1.92^{+0.58}_{-0.51}$  GeV measured using the *t*-channel single top production cross section, and the branching fraction  $\text{br}(t \rightarrow Wb) = 0.962^{+0.068}_{-0.066}(\text{stat})^{+0.064}_{-0.052}(\text{syst})$ . The  $\Gamma(t \rightarrow Wb)$  measurement gives the 95% CL lowerbound of  $\Gamma(t \rightarrow Wb)$  and hence that of  $\Gamma_t$ .
- <sup>5</sup> Results are based on 4.3 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at  $\sqrt{s}$  = 1.96 TeV. The top quark mass and the hadronically decaying *W* boson mass are reconstructed for each candidate events and compared with templates of different top quark width. The two sided 68% CL interval is 0.3 GeV <  $\Gamma_t$  < 4.4 GeV for  $m_t = 172.5$  GeV.
- <sup>6</sup> Based on 955 pb<sup>-1</sup> of  $p\bar{p}$  collision data at  $\sqrt{s}$  = 1.96 TeV. AALTONEN 09M selected *t* $\bar{t}$  candidate events for the  $\ell$  +  $\cancel{E}_T$  + jets channel with one or two *b*-tags, and examine the decay width dependence of the reconstructed  $m_t$  distribution. The result is for  $m_t = 175$  GeV, whereas the upper limit is lower for smaller  $m_t$ .

## t DECAY MODES

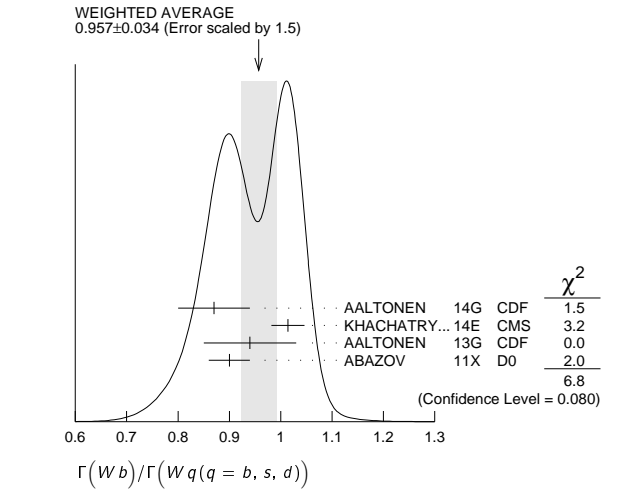
Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $t \rightarrow W q (q = b, s, d)$		
$\Gamma_2$ $t \rightarrow W b$		
$\Gamma_3$ $t \rightarrow \ell \nu_\ell$ anything	[a,b] ( 9.4 $\pm$ 2.4) %	
$\Gamma_4$ $t \rightarrow e \nu_e b$	(13.3 $\pm$ 0.6) %	
$\Gamma_5$ $t \rightarrow \mu \nu_\mu b$	(13.4 $\pm$ 0.6) %	
$\Gamma_6$ $t \rightarrow \tau \nu_\tau b$		
$\Gamma_7$ $t \rightarrow q \bar{q} b$	(66.5 $\pm$ 1.4) %	
$\Gamma_8$ $t \rightarrow \gamma q (q=u,c)$	[c] < 5.9 $\times 10^{-3}$	95%
<b><math>\Delta T = 1</math> weak neutral current (<i>T</i><sub>1</sub>) modes</b>		
$\Gamma_9$ $t \rightarrow Z q (q=u,c)$	$\pi_1$ [d] < 5 $\times 10^{-4}$	95%
$\Gamma_{10}$ $t \rightarrow H q$		
$\Gamma_{11}$ $t \rightarrow \ell^+ \bar{q} \bar{q}' (q=d,s,b; q'=u,c)$	< 1.6 $\times 10^{-3}$	95%

- [a]  $\ell$  means *e* or  $\mu$  decay mode, not the sum over them.
- [b] Assumes lepton universality and *W*-decay acceptance.
- [c] This limit is for  $\Gamma(t \rightarrow \gamma q)/\Gamma(t \rightarrow Wb)$ .
- [d] This limit is for  $\Gamma(t \rightarrow Z q)/\Gamma(t \rightarrow Wb)$ .

## t BRANCHING RATIOS

$\Gamma(Wb)/\Gamma(Wq(q=b,s,d))$			$\Gamma_2/\Gamma_1$
OUR AVERAGE assumes that the systematic uncertainties are uncorrelated.			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.957<math>\pm</math>0.034 OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
0.87 $\pm$ 0.07	<sup>1</sup> AALTONEN	14G CDF	$\ell\ell+E_T \geq 2j$ (0,1,2 b-tag)
1.014 $\pm$ 0.003 $\pm$ 0.032	<sup>2</sup> KHACHATRYAN...14E	CMS	$\ell\ell+E_T + 2,3,4j$ (0-2b-tag)
0.94 $\pm$ 0.09	<sup>3</sup> AALTONEN	13G CDF	$\ell+E_T \geq 3j$ ets ( $\geq 1$ b-tag)
0.90 $\pm$ 0.04	<sup>4</sup> ABAZOV	11X D0	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.97 $\pm$ 0.09 $\pm$ 0.08	<sup>5</sup> ABAZOV	08M D0	$\ell + n$ jets with 0,1,2 b-tag
1.03 $\pm$ 0.19 $\pm$ 0.17	<sup>6</sup> ABAZOV	06K D0	
1.12 $\pm$ 0.21 $\pm$ 0.17 $\pm$ 0.19 $\pm$ 0.13	<sup>7</sup> ACOSTA	05A CDF	Repl. by AALTONEN 13G
0.94 $\pm$ 0.26 $\pm$ 0.17 $\pm$ 0.21 $\pm$ 0.12	<sup>8</sup> AFFOLDER	01c CDF	

- <sup>1</sup> Based on 8.7 fb<sup>-1</sup> of data. This measurement gives  $|V_{tb}| = 0.93 \pm 0.04$  and  $|V_{tb}| > 0.85$  (95% CL) in the SM.
- <sup>2</sup> Based on 19.7 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s}$  = 8 TeV. The result is obtained by counting the number of *b* jets per *t* $\bar{t}$  signal events in the dilepton channel. The *t* $\bar{t}$  production cross section is measured to be  $\sigma(t\bar{t}) = 238 \pm 1 \pm 15$  pb, in good agreement with the SM prediction and the latest CMS measurement of CHATRCHYAN 14F. The measurement gives  $R > 0.995$  (95% CL), or  $|V_{tb}| > 0.975$  (95% CL) in the SM, requiring  $R \leq 1$ .
- <sup>3</sup> Based on 8.7 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s}$  = 1.96 TeV. Measure the fraction of  $t \rightarrow Wb$  decays simultaneously with the *t* $\bar{t}$  cross section. The correlation coefficient between those two measurements is -0.434. Assume unitarity of the 3x3 CKM matrix and set  $|V_{tb}| > 0.89$  at 95% CL.
- <sup>4</sup> Based on 5.4 fb<sup>-1</sup> of data. The error is statistical and systematic combined. The result is a combination of 0.95  $\pm$  0.07 from  $\ell$  + jets channel and 0.86  $\pm$  0.05 from  $\ell\ell$  channel.  $|V_{tb}| = 0.95 \pm 0.02$  follows from the result by assuming unitarity of the 3x3 CKM matrix.
- <sup>5</sup> Result is based on 0.9 fb<sup>-1</sup> of data. The 95% CL lower bound  $R > 0.79$  gives  $|V_{tb}| > 0.89$  (95% CL).
- <sup>6</sup> ABAZOV 06K result is from the analysis of *t* $\bar{t}$   $\rightarrow \ell\nu + \geq 3$  jets with 230 pb<sup>-1</sup> of data at  $\sqrt{s}$  = 1.96 TeV. It gives  $R > 0.61$  and  $|V_{tb}| > 0.78$  at 95% CL. Superseded by ABAZOV 08M.
- <sup>7</sup> ACOSTA 05A result is from the analysis of lepton + jets and di-lepton + jets final states of *t* $\bar{t}$  candidate events with  $\sim 162$  pb<sup>-1</sup> of data at  $\sqrt{s}$  = 1.96 TeV. The first error is statistical and the second systematic. It gives  $R > 0.61$ , or  $|V_{tb}| > 0.78$  at 95% CL.
- <sup>8</sup> AFFOLDER 01c measures the top-quark decay width ratio  $R = \Gamma(Wb)/\Gamma(Wq)$ , where *q* is a *d*, *s*, or *b* quark, by using the number of events with multiple *b*tags. The first error is statistical and the second systematic. A numerical integration of the likelihood function gives  $R > 0.61$  (0.56) at 90% (95%) CL. By assuming three generation unitarity,  $|V_{tb}| = 0.97^{+0.16}_{-0.12}$  or  $|V_{tb}| > 0.78$  (0.75) at 90% (95%) CL is obtained. The result is based on 109 pb<sup>-1</sup> of data at  $\sqrt{s}$  = 1.8 TeV.



$\Gamma(\nu_e \text{ anything})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	$\Gamma_3/\Gamma$
VALUE			
<b>0.094 ± 0.024</b>	1 ABE	98x CDF	

<sup>1</sup>  $\ell$  means  $e$  or  $\mu$  decay mode, not the sum. Assumes lepton universality and  $W$ -decay acceptance.

$\Gamma(\nu_e b)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE				
<b>0.133 ± 0.004 ± 0.005</b>	1 AAD	15cc ATLS	$\ell$ +jets, $\ell\ell$ +jets, $\ell\tau_h$ +jets	

<sup>1</sup> AAD 15cc based on 4.6 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. It is assumed that the top branching ratios to leptons and jets add up to one and that only SM processes contribute to the background. The event selection criteria are optimized for the  $\ell\tau_h$  + jets channel.

$\Gamma(\mu\nu_b)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE				
<b>0.134 ± 0.003 ± 0.005</b>	1 AAD	15cc ATLS	$\ell$ +jets, $\ell\ell$ +jets, $\ell\tau_h$ +jets	

<sup>1</sup> AAD 15cc based on 4.6 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. It is assumed that the top branching ratios to leptons and jets add up to one and that only SM processes contribute to the background. The event selection criteria are optimized for the  $\ell\tau_h$  + jets channel.

$\Gamma(\tau\nu_\tau)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
VALUE				
<b>0.071 ± 0.006 OUR AVERAGE</b>				

0.070 ± 0.003 ± 0.005	1 AAD	15cc ATLS	$\ell$ +jets, $\ell\ell$ +jets, $\ell\tau_h$ +jets
0.096 ± 0.028	2 AALTONEN	14A CDF	$\ell$ + $\tau_h$ + $\geq 2$ jets ( $\geq 1b$ -tag)
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	3 ABULENCIA	06R CDF	$\ell\tau$ + jets
	4 ABE	97v CDF	$\ell\tau$ + jets

<sup>1</sup> AAD 15cc based on 4.6 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. It is assumed that the top branching ratios to leptons and jets add up to one and that only SM processes contribute to the background. The event selection criteria are optimized for the  $\ell\tau_h$  + jets channel.

<sup>2</sup> Based on 9 fb<sup>-1</sup> of data. The measurement is in the channel  $t\bar{t} \rightarrow (b\ell\nu)(\tau\nu)$ , where  $\tau$  decays into hadrons ( $\tau_h$ ), and  $\ell$  ( $e$  or  $\mu$ ) include  $\ell$  from  $\tau$  decays ( $\tau_\ell$ ). The result is consistent with lepton universality.

<sup>3</sup> ABULENCIA 06R looked for  $t\bar{t} \rightarrow (\ell\nu_\ell)(\tau\nu_\tau)b\bar{b}$  events in 194 pb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. 2 events are found where 1.00 ± 0.17 signal and 1.29 ± 0.25 background events are expected, giving a 95% CL upper bound for the partial width ratio  $\Gamma(t \rightarrow \tau\nu)/\Gamma_{SM}(t \rightarrow \tau\nu) < 5.2$ .

<sup>4</sup> ABE 97v searched for  $t\bar{t} \rightarrow (\ell\nu_\ell)(\tau\nu_\tau)b\bar{b}$  events in 109 pb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.8$  TeV. They observed 4 candidate events where one expects  $\sim 1$  signal and  $\sim 2$  background events. Three of the four observed events have jets identified as  $b$  candidates.

$\Gamma(q\bar{q}b)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE				
<b>0.665 ± 0.004 ± 0.013</b>	1 AAD	15cc ATLS	$\ell$ +jets, $\ell\ell$ +jets, $\ell\tau_h$ +jets	

<sup>1</sup> AAD 15cc based on 4.6 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. Branching ratio of top quark into  $b$  and jets. It is assumed that the top branching ratios to leptons and jets add up to one and that only SM processes contribute to the background. The event selection criteria are optimized for the  $\ell\tau_h$  + jets channel.

$\Gamma(\gamma q(q=u,c))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE					
<b>&lt; 0.0059</b>	95	1 CHEKANOV 03	ZEUS	$B(t \rightarrow \gamma u)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.0064	95	2 AARON	09A	$t \rightarrow \gamma u$	
< 0.0465	95	3 ABDALLAH	04C	DLPH $B(\gamma c \text{ or } \gamma u)$	
< 0.0132	95	4 AKTAS	04	H1 $B(t \rightarrow \gamma u)$	
< 0.041	95	5 ACHARD	02J	L3 $B(t \rightarrow \gamma c \text{ or } \gamma u)$	
< 0.032	95	6 ABE	98G	CDF $t\bar{t} \rightarrow (Wb)(\gamma c \text{ or } \gamma u)$	

<sup>1</sup> CHEKANOV 03 looked for single top production via FCNC in the reaction  $e^\pm p \rightarrow e^\pm (t \text{ or } \bar{t}) X$  in 130.1 pb<sup>-1</sup> of data at  $\sqrt{s}=300\text{--}318$  GeV. No evidence for top production and its decay into  $bW$  was found. The result is obtained for  $m_t=175$  GeV when  $B(\gamma c)=B(Zq)=0$ , where  $q$  is a  $u$  or  $c$  quark. Bounds on the effective  $t$ - $u$ - $\gamma$  and  $t$ - $u$ - $Z$  couplings are found in their Fig. 4. The conversion to the constraint listed is from private communication, E. Gallo, January 2004.

<sup>2</sup> AARON 09A looked for single top production via FCNC in  $e^\pm p$  collisions at HERA with 474 pb<sup>-1</sup>. The upper bound of the cross section gives the bound on the FCNC coupling  $\kappa_{t u \gamma}/\Lambda < 1.03$  TeV<sup>-1</sup>, which corresponds to the result for  $m_t = 175$  GeV.

<sup>3</sup> ABDALLAH 04c looked for single top production via FCNC in the reaction  $e^+ e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 541 pb<sup>-1</sup> of data at  $\sqrt{s}=189\text{--}208$  GeV. No deviation from the SM is found, which leads to the bound on  $B(t \rightarrow \gamma q)$ , where  $q$  is a  $u$  or a  $c$  quark, for  $m_t = 175$  GeV when  $B(t \rightarrow Zq)=0$  is assumed. The conversion to the listed bound is from private communication, O. Yushchenko, April 2005. The bounds on the effective  $t$ - $q$ - $\gamma$  and  $t$ - $q$ - $Z$  couplings are given in their Fig. 7 and Table 4, for  $m_t = 170\text{--}180$  GeV, where most conservative bounds are found by choosing the chiral couplings to maximize the negative interference between the virtual  $\gamma$  and  $Z$  exchange amplitudes.

<sup>4</sup> AKTAS 04 looked for single top production via FCNC in  $e^\pm$  collisions at HERA with 118.3 pb<sup>-1</sup>, and found 5 events in the  $e$  or  $\mu$  channels. By assuming that they are due to statistical fluctuation, the upper bound on the  $t u \gamma$  coupling  $\kappa_{t u \gamma} < 0.27$  (95% CL) is obtained. The conversion to the partial width limit, when  $B(\gamma c) = B(Zu) = B(Zc) = 0$ , is from private communication, E. Perez, May 2005.

<sup>5</sup> ACHARD 02J looked for single top production via FCNC in the reaction  $e^+ e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 634 pb<sup>-1</sup> of data at  $\sqrt{s}=189\text{--}209$  GeV. No deviation from the SM is found, which leads to a bound on the top-quark decay branching fraction  $B(\gamma q)$ , where  $q$  is a  $u$  or  $c$  quark. The bound assumes  $B(Zq)=0$  and is for  $m_t = 175$  GeV; bounds for  $m_t=170$  GeV and 180 GeV and  $B(Zq) \neq 0$  are given in Fig. 5 and Table 7.

<sup>6</sup> ABE 98G looked for  $t\bar{t}$  events where one  $t$  decays into  $q\gamma$  while the other decays into  $bW$ . The quoted bound is for  $\Gamma(\gamma q)/\Gamma(Wb)$ .

$\Gamma(Zq(q=u,c))/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
Test for $\Delta T=1$ weak neutral current. Allowed by higher-order electroweak interaction.	

< 0.7	95	1 AAD	16D ATLS	$t \rightarrow Zq (q = u, c)$
< 0.5	95	2 CHATRCHYAN14s	CMS	$t \rightarrow Zq (q = u, c)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.6	95	3 CHATRCHYAN14s	CMS	$t \rightarrow Zq (q = u, c)$
< 2.1	95	4 CHATRCHYAN13F	CMS	$t \rightarrow Zq (q = u, c)$
< 7.3	95	5 AAD	12BT ATLS	$t\bar{t} \rightarrow \ell^+ \ell^- \ell^\pm + \bar{b}T + \text{jets}$
< 32	95	6 ABAZOV	11M D0	$t \rightarrow Zq (q = u, c)$
< 83	95	7 AALTONEN	09AL CDF	$t \rightarrow Zq (q = c)$
< 37	95	8 AALTONEN	08AD CDF	$t \rightarrow Zq (q = u, c)$
< 1.59 × 10 <sup>2</sup>	95	9 ABDALLAH	04C DLPH	$e^+ e^- \rightarrow \bar{t}c$ or $\bar{t}u$
< 1.37 × 10 <sup>2</sup>	95	10 ACHARD	02J L3	$e^+ e^- \rightarrow \bar{t}c$ or $\bar{t}u$
< 1.4 × 10 <sup>2</sup>	95	11 HEISTER	02Q ALEP	$e^+ e^- \rightarrow \bar{t}c$ or $\bar{t}u$
< 1.37 × 10 <sup>2</sup>	95	12 ABBIENDI	01T OPAL	$e^+ e^- \rightarrow \bar{t}c$ or $\bar{t}u$
< 1.7 × 10 <sup>2</sup>	95	13 BARATE	00S ALEP	$e^+ e^- \rightarrow \bar{t}c$ or $\bar{t}u$
< 3.3 × 10 <sup>2</sup>	95	14 ABE	98G CDF	$t\bar{t} \rightarrow (Wb)(Zc \text{ or } Zu)$

<sup>1</sup> AAD 16D based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. The FCNC decay is searched for in  $t\bar{t}$  events in the final state  $(bW)(qZ)$  when both  $W$  and  $Z$  decay leptonically, giving 3 charged leptons.

<sup>2</sup> CHATRCHYAN 14s combined search limit from this and CHATRCHYAN 13F data.

<sup>3</sup> Based on 19.7 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. The flavor changing decay is searched for in  $t\bar{t}$  events in the final state  $(bW)(qZ)$  when both  $W$  and  $Z$  decay leptonically, giving 3 charged leptons.

<sup>4</sup> Based on 5.0 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. Search for FCNC decays of the top quark in  $t\bar{t} \rightarrow \ell^+ \ell^- \ell^\pm \nu$  + jets ( $\ell, \ell' = e, \mu$ ) final states found no excess of signal events.

<sup>5</sup> Based on 2.1 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV.

<sup>6</sup> Based on 4.1 fb<sup>-1</sup> of data. ABAZOV 11M searched for FCNC decays of the top quark in  $t\bar{t} \rightarrow \ell^+ \ell^- \ell^\pm \nu$  + jets ( $\ell, \ell' = e, \mu$ ) final states, and absence of the signal gives the bound.

<sup>7</sup> Based on  $p\bar{p}$  data of 1.52 fb<sup>-1</sup>. AALTONEN 09AL compared  $t\bar{t} \rightarrow WbWb \rightarrow \ell\nu b j j b$  and  $t\bar{t} \rightarrow ZcWb \rightarrow \ell\ell c j j b$  decay chains, and absence of the latter signal gives the bound. The result is for 100% longitudinally polarized  $Z$  boson and the theoretical  $t\bar{t}$  production cross section. The results for different  $Z$  polarizations and those without the cross section assumption are given in their Table XII.

<sup>8</sup> Result is based on 1.9 fb<sup>-1</sup> of data at  $\sqrt{s} = 1.96$  TeV.  $t\bar{t} \rightarrow WbZq$  or  $ZqZq$  processes have been looked for in  $Z + \geq 4$  jet events with and without  $b$ -tag. No signal leads to the bound  $B(t \rightarrow Zq) < 0.037$  (0.041) for  $m_t = 175$  (170) GeV.

<sup>9</sup> ABDALLAH 04c looked for single top production via FCNC in the reaction  $e^+ e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 541 pb<sup>-1</sup> of data at  $\sqrt{s}=189\text{--}208$  GeV. No deviation from the SM is found, which leads to the bound on  $B(t \rightarrow \gamma q)$ , where  $q$  is a  $u$  or a  $c$  quark, for  $m_t = 175$  GeV when  $B(t \rightarrow Zq)=0$  is assumed. The conversion to the listed bound is from private communication, O. Yushchenko, April 2005. The bounds on the effective  $t$ - $q$ - $\gamma$  and  $t$ - $q$ - $Z$  couplings are given in their Fig. 7 and Table 4, for  $m_t = 170\text{--}180$  GeV, where most conservative bounds are found by choosing the chiral couplings to maximize the negative interference between the virtual  $\gamma$  and  $Z$  exchange amplitudes.

<sup>10</sup> ACHARD 02J looked for single top production via FCNC in the reaction  $e^+ e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 634 pb<sup>-1</sup> of data at  $\sqrt{s}=189\text{--}209$  GeV. No deviation from the SM is found, which leads to a bound on the top-quark decay branching fraction  $B(Zq)$ , where  $q$  is a  $u$  or  $c$  quark. The bound assumes  $B(\gamma q)=0$  and is for  $m_t = 175$  GeV; bounds for  $m_t=170$  GeV and 180 GeV and  $B(\gamma q) \neq 0$  are given in Fig. 5 and Table 7. Table 6 gives constraints on  $t$ - $c$ - $e$ - $e$  four-fermi contact interactions.

<sup>11</sup> HEISTER 02Q looked for single top production via FCNC in the reaction  $e^+ e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 214 pb<sup>-1</sup> of data at  $\sqrt{s}=204\text{--}209$  GeV. No deviation from the SM is found, which leads to a bound on the branching fraction  $B(Zq)$ , where  $q$  is a  $u$  or  $c$  quark. The bound assumes  $B(\gamma q)=0$  and is for  $m_t = 174$  GeV. Bounds on the effective  $t$ - $(c \text{ or } u)$ - $\gamma$  and  $t$ - $(c \text{ or } u)$ - $Z$  couplings are given in their Fig. 2.

<sup>12</sup> ABBIENDI 01T looked for single top production via FCNC in the reaction  $e^+ e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 600 pb<sup>-1</sup> of data at  $\sqrt{s}=189\text{--}209$  GeV. No deviation from the SM is found, which leads to bounds on the branching fractions  $B(Zq)$  and  $B(\gamma q)$ , where  $q$  is a  $u$

# Quark Particle Listings

*t*

or *c* quark. The result is obtained for  $m_t = 174$  GeV. The upper bound becomes 9.7% (20.6%) for  $m_t = 169$  (179) GeV. Bounds on the effective  $t$ -(*c* or *u*)- $\gamma$  and  $t$ -(*c* or *u*)- $Z$  couplings are given in their Fig. 4.

- <sup>13</sup> BARATE 00s looked for single top production via FCNC in the reaction  $e^+e^- \rightarrow \bar{t}c$  or  $\bar{t}u$  in 411 pb<sup>-1</sup> of data at c.m. energies between 189 and 202 GeV. No deviation from the SM is found, which leads to a bound on the branching fraction. The bound assumes  $B(\gamma q)=0$ . Bounds on the effective  $t$ -(*c* or *u*)- $\gamma$  and  $t$ -(*c* or *u*)- $Z$  couplings are given in their Fig. 4.

- <sup>14</sup> ABE 98G looked for  $t\bar{t}$  events where one *t* decays into three jets and the other decays into  $qZ$  with  $Z \rightarrow \ell\ell$ . The quoted bound is for  $\Gamma(Zq)/\Gamma(Wb)$ .

$\Gamma(Hq)/\Gamma_{\text{total}}$		$\Gamma_{10}/\Gamma$		
VALUE (units 10 <sup>-3</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
< 5.6	95	<sup>1</sup> AAD	15co ATLS	$t \rightarrow Hc(H \rightarrow bb)$
< 6.1	95	<sup>1</sup> AAD	15co ATLS	$t \rightarrow Hu(H \rightarrow bb)$
< 5.6	95	<sup>2</sup> KHACHATRYAN14Q	CMS	$t \rightarrow Hc(H \rightarrow \gamma\gamma \text{ or lepton})$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< 7.9	95	<sup>3</sup> AAD	14AA ATLS	$t \rightarrow Hq(q=u,c; H \rightarrow \gamma\gamma)$
< 13	95	<sup>4</sup> CHATRCHYAN14R	CMS	$t \rightarrow Hc(H \rightarrow \geq 2\ell)$

- <sup>1</sup> AAD 15co based on 20.3 fb<sup>-1</sup> at  $\sqrt{s} = 8$  TeV of *pp* data. Searches for  $t\bar{t}$  events, where the other top quark decays semi-leptonically. Exploits high multiplicity of *b*-jets and uses a likelihood discriminant. Combining with other ATLAS searches for different Higgs decay modes,  $B(t \rightarrow Hc) < 0.46\%$  and  $B(t \rightarrow Hu) < 0.45\%$  are obtained.

- <sup>2</sup> KHACHATRYAN 14Q based on 19.5 fb<sup>-1</sup> at  $\sqrt{s} = 8$  TeV of *pp* data. Search for final states with  $\geq 3$  isolated charged leptons or with a photon pair accompanied by  $\geq 1$  lepton(s).

- <sup>3</sup> AAD 14AA based on 4.7 fb<sup>-1</sup> at  $\sqrt{s} = 7$  TeV and 20.3 fb<sup>-1</sup> at  $\sqrt{s} = 8$  TeV of *pp* data. The upper-bound is for the sum of  $\text{Br}(t \rightarrow Hc)$  and  $\text{Br}(t \rightarrow Hu)$ . Search for  $t\bar{t}$  events, where the other top quark decays hadronically or semi-leptonically. The upper bound constrains the *H*-*t*-*c* Yukawa couplings  $\sqrt{|Y_{tCl}^H|^2 + |Y_{tCR}^H|^2} < 0.17$  (95% CL).

- <sup>4</sup> Based on 19.5 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 8$  TeV. Search for final states with 3 or more isolated high  $E_T$  charged leptons ( $\ell = e, \mu$ ) bounds the  $t \rightarrow Hc$  decay in  $t\bar{t}$  events when *H* decays contain a pair of leptons. The upper bound constrains the *H*-*t*-*c* Yukawa couplings  $\sqrt{|Y_{tCl}^H|^2 + |Y_{tCR}^H|^2} < 0.21$  (95% CL).

$\Gamma(\ell^+q\bar{q}'(q=d,s,b; q'=u,c))/\Gamma_{\text{total}}$		$\Gamma_{11}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b>1.6 × 10<sup>-3</sup></b>	95	<sup>1</sup> CHATRCHYAN14Q	CMS	$\mu$ + dijets
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.7 × 10 <sup>-3</sup>	95	<sup>1</sup> CHATRCHYAN14Q	CMS	$e$ + dijets

- <sup>1</sup> Based on 19.5 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 8$  TeV. Baryon number violating decays of the top quark are searched for in  $t\bar{t}$  production events where one of the pair decays into hadronic three jets.

## *t*-quark EW Couplings

*W* helicity fractions in top decays.  $F_0$  is the fraction of longitudinal and  $F_+$  the fraction of right-handed *W* bosons.  $F_{V+A}$  is the fraction of *V*+*A* current in top decays. The effective Lagrangian (cited by ABZOV 08A) has terms  $t_L^L$  and  $t_R^R$  for *V*-*A* and *V*+*A* couplings,  $t_L^T$  and  $t_R^T$  for tensor couplings with  $b_R$  and  $b_L$  respectively.

$F_0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.690 ± 0.030 OUR AVERAGE</b>			
0.726 ± 0.066 ± 0.067	<sup>1</sup> AALTONEN	13D CDF	$F_0 = B(t \rightarrow W_0b)$
0.682 ± 0.030 ± 0.033	<sup>2</sup> CHATRCHYAN13BH	CMS	$F_0 = B(t \rightarrow W_0b)$
0.67 ± 0.07	<sup>3</sup> AAD	12Bg ATLS	$F_0 = B(t \rightarrow W_0b)$
0.722 ± 0.062 ± 0.052	<sup>4</sup> AALTONEN	12Z TEVA	$F_0 = B(t \rightarrow W_0b)$
0.669 ± 0.078 ± 0.065	<sup>5</sup> ABZOV	11c D0	$F_0 = B(t \rightarrow W_0b)$
0.91 ± 0.37 ± 0.13	<sup>6</sup> AFFOLDER	00B CDF	$F_0 = B(t \rightarrow W_0b)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.70 ± 0.07 ± 0.04	<sup>7</sup> AALTONEN	10Q CDF	Repl. by AALTONEN 12Z
0.62 ± 0.10 ± 0.05	<sup>8</sup> AALTONEN	09Q CDF	Repl. by AALTONEN 10Q
0.425 ± 0.166 ± 0.102	<sup>9</sup> ABZOV	08B D0	Repl. by ABZOV 11c
0.85 <sup>+0.15</sup> <sub>-0.22</sub> ± 0.06	<sup>10</sup> ABULENCIA	07I CDF	$F_0 = B(t \rightarrow W_0b)$
0.74 <sup>+0.22</sup> <sub>-0.34</sub>	<sup>11</sup> ABULENCIA	06U CDF	$F_0 = B(t \rightarrow W_0b)$
0.56 ± 0.31	<sup>12</sup> ABZOV	05G D0	$F_0 = B(t \rightarrow W_0b)$

- <sup>1</sup> Based on 8.7 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV using  $t\bar{t}$  events with  $\ell + \cancel{E}_T + \geq 4$  jets ( $\geq 1b$ ), and under the constraint  $F_0 + F_+ + F_- = 1$ . The statistical errors of  $F_0$  and  $F_+$  are correlated with correlation coefficient  $\rho(F_0, F_+) = -0.69$ .

- <sup>2</sup> Based on 5.0 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 13BH studied *tt* events with large  $\cancel{E}_T$  and  $\ell + \geq 4$  jets using a constrained kinematic fit.

- <sup>3</sup> Based on 1.04 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 7$  TeV. AAD 12Bg studied *tt* events with large  $\cancel{E}_T$  and either  $\ell + \geq 4j$  or  $\ell\ell + \geq 2j$ . The uncertainties are not independent,  $\rho(F_0, F_-) = -0.96$ .

- <sup>4</sup> Based on 2.7 and 5.1 fb<sup>-1</sup> of CDF data in  $\ell$  + jets and dilepton channels, and 5.4 fb<sup>-1</sup> of D0 data in  $\ell$  + jets and dilepton channels.  $F_0 = 0.682 \pm 0.035 \pm 0.046$  if  $F_+ = 0.0017(1)$ , while  $F_+ = -0.015 \pm 0.018 \pm 0.030$  if  $F_0 = 0.688(4)$ , where the assumed fixed values are the SM prediction for  $m_t = 173.3 \pm 1.1$  GeV and  $m_W = 80.399 \pm 0.023$  GeV.

- <sup>5</sup> Results are based on 5.4 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at 1.96 TeV, including those of ABZOV 08B. Under the SM constraint of  $f_0 = 0.698$  (for  $m_t = 173.3$  GeV,  $m_W = 80.399$  GeV),  $f_+ = 0.010 \pm 0.022 \pm 0.030$  is obtained.

- <sup>6</sup> AFFOLDER 00B studied the angular distribution of leptonic decays of *W* bosons in  $t \rightarrow Wb$  events. The ratio  $F_0$  is the fraction of the helicity zero (longitudinal) *W* bosons in the decaying top quark rest frame.  $B(t \rightarrow W_+b)$  is the fraction of positive helicity (right-handed) positive charge *W* bosons in the top quark decays. It is obtained by assuming the Standard Model value of  $F_0$ .

- <sup>7</sup> Results are based on 2.7 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.  $F_0$  result is obtained by assuming  $F_+ = 0$ , while  $F_+$  result is obtained for  $F_0 = 0.70$ , the SM value. Model independent fits for the two fractions give  $F_0 = 0.88 \pm 0.11 \pm 0.06$  and  $F_+ = -0.15 \pm 0.07 \pm 0.06$  with correlation coefficient of  $-0.59$ . The results are for  $m_t = 175$  GeV.

- <sup>8</sup> Results are based on 1.9 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.  $F_0$  result is obtained assuming  $F_+ = 0$ , while  $F_+$  result is obtained for  $F_0 = 0.70$ , the SM values. Model independent fits for the two fractions give  $F_0 = 0.66 \pm 0.16 \pm 0.05$  and  $F_+ = -0.03 \pm 0.06 \pm 0.03$ .

- <sup>9</sup> Based on 1 fb<sup>-1</sup> at  $\sqrt{s} = 1.96$  TeV.

- <sup>10</sup> Based on 318 pb<sup>-1</sup> of data at  $\sqrt{s} = 1.96$  TeV.

- <sup>11</sup> Based on 200 pb<sup>-1</sup> of data at  $\sqrt{s} = 1.96$  TeV.  $t \rightarrow Wb \rightarrow \ell\nu b$  ( $\ell = e$  or  $\mu$ ). The errors are stat + syst.

- <sup>12</sup> ABZOV 05G studied the angular distribution of leptonic decays of *W* bosons in  $t\bar{t}$  candidate events with lepton + jets final states, and obtained the fraction of longitudinally polarized *W* under the constraint of no right-handed current,  $F_+ = 0$ . Based on 125 pb<sup>-1</sup> of data at  $\sqrt{s} = 1.8$  TeV.

$F_-$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.314 ± 0.025 OUR AVERAGE</b>			
0.310 ± 0.022 ± 0.022	<sup>1</sup> CHATRCHYAN13BH	CMS	$F_- = B(t \rightarrow W_-b)$
0.32 ± 0.04	<sup>2</sup> AAD	12Bg ATLS	$F_- = B(t \rightarrow W_-b)$

- <sup>1</sup> Based on 5.0 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 13BH studied *tt* events with large  $\cancel{E}_T$  and  $\ell + \geq 4$  jets using a constrained kinematic fit.

- <sup>2</sup> Based on 1.04 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 7$  TeV. AAD 12Bg studied *tt* events with large  $\cancel{E}_T$  and either  $\ell + \geq 4j$  or  $\ell\ell + \geq 2j$ . The uncertainties are not independent,  $\rho(F_0, F_-) = -0.96$ .

$F_+$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.008 ± 0.016 OUR AVERAGE</b>				
-0.045 ± 0.044 ± 0.058		<sup>1</sup> AALTONEN	13D CDF	$F_+ = B(t \rightarrow W_+b)$
0.008 ± 0.012 ± 0.014		<sup>2</sup> CHATRCHYAN13BH	CMS	$F_+ = B(t \rightarrow W_+b)$
0.01 ± 0.05		<sup>3</sup> AAD	12Bg ATLS	$F_+ = B(t \rightarrow W_+b)$
0.023 ± 0.041 ± 0.034		<sup>4</sup> ABZOV	11c D0	$F_+ = B(t \rightarrow W_+b)$
0.11 ± 0.15		<sup>5</sup> AFFOLDER	00B CDF	$F_+ = B(t \rightarrow W_+b)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.033 ± 0.034 ± 0.031		<sup>6</sup> AALTONEN	12Z TEVA	$F_+ = B(t \rightarrow W_+b)$
-0.01 ± 0.02 ± 0.05		<sup>7</sup> AALTONEN	10Q CDF	Repl. by AALTONEN 13D
-0.04 ± 0.04 ± 0.03		<sup>8</sup> AALTONEN	09Q CDF	Repl. by AALTONEN 10Q
0.119 ± 0.090 ± 0.053		<sup>9</sup> ABZOV	08B D0	Repl. by ABZOV 11c
0.056 ± 0.080 ± 0.057		<sup>10</sup> ABZOV	07D D0	$F_+ = B(t \rightarrow W_+b)$
0.05 <sup>+0.11</sup> <sub>-0.05</sub> ± 0.03		<sup>11</sup> ABULENCIA	07I CDF	$F_+ = B(t \rightarrow W_+b)$
< 0.26	95	<sup>11</sup> ABULENCIA	07I CDF	$F_+ = B(t \rightarrow W_+b)$
< 0.27	95	<sup>12</sup> ABULENCIA	06U CDF	$F_+ = B(t \rightarrow W_+b)$
0.00 ± 0.13 ± 0.07		<sup>13</sup> ABZOV	05L D0	$F_+ = B(t \rightarrow W_+b)$
< 0.25	95	<sup>13</sup> ABZOV	05L D0	$F_+ = B(t \rightarrow W_+b)$
< 0.24	95	<sup>14</sup> ACOSTA	05D CDF	$F_+ = B(t \rightarrow W_+b)$

- <sup>1</sup> Based on 8.7 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV using  $t\bar{t}$  events with  $\ell + \cancel{E}_T + \geq 4$  jets ( $\geq 1b$ ), and under the constraint  $F_0 + F_+ + F_- = 1$ . The statistical errors of  $F_0$  and  $F_+$  are correlated with correlation coefficient  $\rho(F_0, F_+) = -0.69$ .

- <sup>2</sup> Based on 5.0 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 13BH studied *tt* events with large  $\cancel{E}_T$  and  $\ell + \geq 4$  jets using a constrained kinematic fit.

- <sup>3</sup> Based on 1.04 fb<sup>-1</sup> of *pp* data at  $\sqrt{s} = 7$  TeV. AAD 12Bg studied *tt* events with large  $\cancel{E}_T$  and either  $\ell + \geq 4j$  or  $\ell\ell + \geq 2j$ .

- <sup>4</sup> Results are based on 5.4 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at 1.96 TeV, including those of ABZOV 08B. Under the SM constraint of  $f_0 = 0.698$  (for  $m_t = 173.3$  GeV,  $m_W = 80.399$  GeV),  $f_+ = 0.010 \pm 0.022 \pm 0.030$  is obtained.

- <sup>5</sup> AFFOLDER 00B studied the angular distribution of leptonic decays of *W* bosons in  $t \rightarrow Wb$  events. The ratio  $F_0$  is the fraction of the helicity zero (longitudinal) *W* bosons in the decaying top quark rest frame.  $B(t \rightarrow W_+b)$  is the fraction of positive helicity (right-handed) positive charge *W* bosons in the top quark decays. It is obtained by assuming the Standard Model value of  $F_0$ .

- <sup>6</sup> Based on 2.7 and 5.1 fb<sup>-1</sup> of CDF data in  $\ell$  + jets and dilepton channels, and 5.4 fb<sup>-1</sup> of D0 data in  $\ell$  + jets and dilepton channels.  $F_0 = 0.682 \pm 0.035 \pm 0.046$  if  $F_+ = 0.0017(1)$ , while  $F_+ = -0.015 \pm 0.018 \pm 0.030$  if  $F_0 = 0.688(4)$ , where the assumed fixed values are the SM prediction for  $m_t = 173.3 \pm 1.1$  GeV and  $m_W = 80.399 \pm 0.023$  GeV.

- <sup>7</sup> Results are based on 2.7 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.  $F_0$  result is obtained by assuming  $F_+ = 0$ , while  $F_+$  result is obtained for  $F_0 = 0.70$ , the SM value. Model independent fits for the two fractions give  $F_0 = 0.88 \pm 0.11 \pm 0.06$  and  $F_+ = -0.15 \pm 0.07 \pm 0.06$  with correlation coefficient of  $-0.59$ . The results are for  $m_t = 175$  GeV.



- <sup>8</sup> Results are based on  $1.9 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .  $F_0$  result is obtained assuming  $F_{\perp} = 0$ , while  $F_{\perp}$  result is obtained for  $F_0 = 0.70$ , the SM values. Model independent fits for the two fractions give  $F_0 = 0.66 \pm 0.16 \pm 0.05$  and  $F_{\perp} = -0.03 \pm 0.06 \pm 0.03$ .
- <sup>9</sup> Based on  $1 \text{ fb}^{-1}$  at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>10</sup> Based on  $370 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ , using the  $\ell + \text{jets}$  and dilepton decay channels. The result assumes  $F_0 = 0.70$ , and it gives  $F_{\perp} < 0.23$  at 95% CL.
- <sup>11</sup> Based on  $318 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>12</sup> Based on  $200 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .  $t \rightarrow Wb \rightarrow \ell\nu b$  ( $\ell = e$  or  $\mu$ ). The errors are stat + syst.
- <sup>13</sup> ABAZOV 05L studied the angular distribution of leptonic decays of  $W$  bosons in  $t\bar{t}$  events, where one of the  $W$ 's from  $t$  or  $\bar{t}$  decays into  $e$  or  $\mu$  and the other decays hadronically. The fraction of the "+" helicity  $W$  boson is obtained by assuming  $F_0 = 0.7$ , which is the generic prediction for any linear combination of V and A currents. Based on  $230 \pm 15 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>14</sup> ACOSTA 05D measures the  $m_{\ell\ell}^2$  distribution in  $t\bar{t}$  production events where one or both  $W$ 's decay leptonically to  $\ell = e$  or  $\mu$ , and finds a bound on the V+A coupling of the  $tb$  vertex. By assuming the SM value of the longitudinal  $W$  fraction  $F_0 = B(t \rightarrow W_0 b) = 0.70$ , the bound on  $F_{\perp}$  is obtained. If the results are combined with those of AFFOLDER 00B, the bounds become  $F_{V+A} < 0.61$  (95% CL) and  $F_{\perp} < 0.18$  (95% CL), respectively. Based on  $109 \pm 7 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.8 \text{ TeV}$  (run I).

 **$F_{V+A}$** 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.29</b>	95	<sup>1</sup> ABULENCIA	07G CDF	$F_{V+A} = B(t \rightarrow Wb_R)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.06 \pm 0.22 \pm 0.12$		<sup>1</sup> ABULENCIA	07G CDF	$F_{V+A} = B(t \rightarrow Wb_R)$
<b>&lt; 0.80</b>	95	<sup>2</sup> ACOSTA	05D CDF	$F_{V+A} = B(t \rightarrow Wb_R)$

- <sup>1</sup> Based on  $700 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ .
- <sup>2</sup> ACOSTA 05D measures the  $m_{\ell\ell}^2$  distribution in  $t\bar{t}$  production events where one or both  $W$ 's decay leptonically to  $\ell = e$  or  $\mu$ , and finds a bound on the V+A coupling of the  $tb$  vertex. By assuming the SM value of the longitudinal  $W$  fraction  $F_0 = B(t \rightarrow W_0 b) = 0.70$ , the bound on  $F_{\perp}$  is obtained. If the results are combined with those of AFFOLDER 00B, the bounds become  $F_{V+A} < 0.61$  (95% CL) and  $F_{\perp} < 0.18$  (95% CL), respectively. Based on  $109 \pm 7 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.8 \text{ TeV}$  (run I).

 **$f_1^R$** 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.20 < \text{Re}(V_{tb} f_1^R) < 0.23$	95	<sup>1</sup> AAD	12BG ATLS	Constr. on $Wtb$ vtx
$(V_{tb} f_1^R)^2 < 0.93$	95	<sup>2</sup> ABAZOV	12E D0	Single-top
$ f_1^R ^2 < 0.30$	95	<sup>3</sup> ABAZOV	12I D0	single- $t + W$ helicity
$ f_1^R ^2 < 1.01$	95	<sup>4</sup> ABAZOV	09J D0	$ f_1^L  = 1,  f_2^L  =  f_2^R  = 0$
$ f_1^R ^2 < 2.5$	95	<sup>5</sup> ABAZOV	08AI D0	$ f_1^L  = 1.8^{+1.0}_{-1.3}$

- <sup>1</sup> Based on  $1.04 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . AAD 12BG studied  $tt$  events with large  $E_T$  and either  $\ell + \geq 4j$  or  $\ell\ell + \geq 2j$ .
- <sup>2</sup> Based on  $5.4 \text{ fb}^{-1}$  of data. For each value of the form factor quoted the other two are assumed to have their SM value. Their Fig. 4 shows two-dimensional posterior probability density distributions for the anomalous couplings.
- <sup>3</sup> Based on  $5.4 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ . Results are obtained by combining the limits from the  $W$  helicity measurements and those from the single top quark production.
- <sup>4</sup> Based on  $1 \text{ fb}^{-1}$  of data at  $p\bar{p}$  collisions  $\sqrt{s} = 1.96 \text{ TeV}$ . Combined result of the  $W$  helicity measurement in  $t\bar{t}$  events (ABAZOV 08B) and the search for anomalous  $tb$   $W$  couplings in the single top production (ABAZOV 08AI). Constraints when  $f_1^L$  and one of the anomalous couplings are simultaneously allowed to vary are given in their Fig. 1 and Table 1.
- <sup>5</sup> Result is based on  $0.9 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Single top quark production events are used to measure the Lorentz structure of the  $tb$   $W$  coupling. The upper bounds on the non-standard couplings are obtained when only one non-standard coupling is allowed to be present together with the SM one,  $f_1^L = V_{tb}^*$ .

 **$f_2^L$** 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.14 < \text{Re}(f_2^L) < 0.11$	95	<sup>1</sup> AAD	12BG ATLS	Constr. on $Wtb$ vtx
$(V_{tb} f_2^L)^2 < 0.13$	95	<sup>2</sup> ABAZOV	12E D0	Single-top
$ f_2^L ^2 < 0.05$	95	<sup>3</sup> ABAZOV	12I D0	single- $t + W$ helicity
$ f_2^L ^2 < 0.28$	95	<sup>4</sup> ABAZOV	09J D0	$ f_1^L  = 1,  f_1^R  =  f_2^R  = 0$
$ f_2^L ^2 < 0.5$	95	<sup>5</sup> ABAZOV	08AI D0	$ f_1^L  = 1.4^{+0.6}_{-0.5}$

- <sup>1</sup> Based on  $1.04 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . AAD 12BG studied  $tt$  events with large  $E_T$  and either  $\ell + \geq 4j$  or  $\ell\ell + \geq 2j$ .
- <sup>2</sup> Based on  $5.4 \text{ fb}^{-1}$  of data. For each value of the form factor quoted the other two are assumed to have their SM value. Their Fig. 4 shows two-dimensional posterior probability density distributions for the anomalous couplings.
- <sup>3</sup> Based on  $5.4 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ . Results are obtained by combining the limits from the  $W$  helicity measurements and those from the single top quark production.
- <sup>4</sup> Based on  $1 \text{ fb}^{-1}$  of data at  $p\bar{p}$  collisions  $\sqrt{s} = 1.96 \text{ TeV}$ . Combined result of the  $W$  helicity measurement in  $t\bar{t}$  events (ABAZOV 08B) and the search for anomalous  $tb$   $W$  couplings in the single top production (ABAZOV 08AI). Constraints when  $f_1^L$  and one of the anomalous couplings are simultaneously allowed to vary are given in their Fig. 1 and Table 1.
- <sup>5</sup> Result is based on  $0.9 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Single top quark production events are used to measure the Lorentz structure of the  $tb$   $W$  coupling. The upper bounds on the non-standard couplings are obtained when only one non-standard coupling is allowed to be present together with the SM one,  $f_1^L = V_{tb}^*$ .

 **$f_2^R$** 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.08 < \text{Re}(f_2^R) < 0.04$	95	<sup>1</sup> AAD	12BG ATLS	Constr. on $Wtb$ vtx
$(V_{tb} f_2^R)^2 < 0.06$	95	<sup>2</sup> ABAZOV	12E D0	Single-top
$ f_2^R ^2 < 0.12$	95	<sup>3</sup> ABAZOV	12I D0	single- $t + W$ helicity
$ f_2^R ^2 < 0.23$	95	<sup>4</sup> ABAZOV	09J D0	$ f_1^L  = 1,  f_1^R  =  f_2^L  = 0$
$ f_2^R ^2 < 0.3$	95	<sup>5</sup> ABAZOV	08AI D0	$ f_1^L  = 1.4^{+0.9}_{-0.8}$

- <sup>1</sup> Based on  $1.04 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . AAD 12BG studied  $tt$  events with large  $E_T$  and either  $\ell + \geq 4j$  or  $\ell\ell + \geq 2j$ .
- <sup>2</sup> Based on  $5.4 \text{ fb}^{-1}$  of data. For each value of the form factor quoted the other two are assumed to have their SM value. Their Fig. 4 shows two-dimensional posterior probability density distributions for the anomalous couplings.
- <sup>3</sup> Based on  $5.4 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at  $1.96 \text{ TeV}$ . Results are obtained by combining the limits from the  $W$  helicity measurements and those from the single top quark production.
- <sup>4</sup> Based on  $1 \text{ fb}^{-1}$  of data at  $p\bar{p}$  collisions  $\sqrt{s} = 1.96 \text{ TeV}$ . Combined result of the  $W$  helicity measurement in  $t\bar{t}$  events (ABAZOV 08B) and the search for anomalous  $tb$   $W$  couplings in the single top production (ABAZOV 08AI). Constraints when  $f_1^L$  and one of the anomalous couplings are simultaneously allowed to vary are given in their Fig. 1 and Table 1.
- <sup>5</sup> Result is based on  $0.9 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . Single top quark production events are used to measure the Lorentz structure of the  $tb$   $W$  coupling. The upper bounds on the non-standard couplings are obtained when only one non-standard coupling is allowed to be present together with the SM one,  $f_1^L = V_{tb}^*$ .

**Spin Correlation in  $t\bar{t}$  Production in  $p\bar{p}$  Collisions**

$C$  is the correlation strength parameter,  $f$  is the ratio of events with correlated  $t$  and  $\bar{t}$  spins (SM prediction:  $f = 1$ ), and  $\kappa$  is the spin correlation coefficient. See "The Top Quark" review for more information.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.85 \pm 0.29$	<sup>1</sup> ABAZOV	12B D0	$f(\ell\ell + \geq 2 \text{ jets}, \ell + \geq 4 \text{ jets})$
$1.15^{+0.42}_{-0.43}$	<sup>2</sup> ABAZOV	12B D0	$f(\ell + E_T + \geq 4 \text{ jets})$
$0.60^{+0.50}_{-0.16}$	<sup>3</sup> AALTONEN	11AR CDF	$\kappa(\ell + E_T + \geq 4 \text{ jets})$
$0.74^{+0.40}_{-0.41}$	<sup>4</sup> ABAZOV	11AE D0	$f(\ell\ell + E_T + \geq 2 \text{ jets})$
$0.10 \pm 0.45$	<sup>5</sup> ABAZOV	11AF D0	$C(\ell\ell + E_T + \geq 2 \text{ jets})$

- <sup>1</sup> This is a combination of the lepton + jets analysis presented in ABAZOV 12B and the dilepton measurement of ABAZOV 11AE. It provides a  $3.1 \sigma$  evidence for the  $t\bar{t}$  spin correlation.
- <sup>2</sup> Based on  $5.3 \text{ fb}^{-1}$  of data. The error is statistical and systematic combined. A matrix element method is used.
- <sup>3</sup> Based on  $4.3 \text{ fb}^{-1}$  of data. The measurement is based on the angular study of the top quark decay products in the helicity basis. The theory prediction is  $\kappa \approx 0.40$ .
- <sup>4</sup> Based on  $5.4 \text{ fb}^{-1}$  of data using a matrix element method. The error is statistical and systematic combined. The no-correlation hypothesis is excluded at the 97.7% CL.
- <sup>5</sup> Based on  $5.4 \text{ fb}^{-1}$  of data. The error is statistical and systematic combined. The NLO QCD prediction is  $C = 0.78 \pm 0.03$ . The neutrino weighting method is used for reconstruction of kinematics.

**Spin Correlation in  $t\bar{t}$  Production in  $pp$  Collisions**

Spin correlation,  $f_{SM}$ , measures the strength of the correlation between the spins of the pair produced  $t\bar{t}$ .  $f_{SM} = 1$  for the SM, while  $f_{SM} = 0$  for no spin correlation.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.20 \pm 0.05 \pm 0.13$	<sup>1</sup> AAD	15J ATLS	$\Delta\phi(\ell\ell)$ in $\ell\ell + \geq 2j$ ( $\geq 1b$ )
$1.19 \pm 0.09 \pm 0.18$	<sup>2</sup> AAD	14BB ATLS	$\Delta\phi(\ell\ell)$ in $\ell\ell + \geq 2j$ events
$1.12 \pm 0.11 \pm 0.22$	<sup>2</sup> AAD	14BB ATLS	$\Delta\phi(\ell j)$ in $\ell + \geq 4j$ events
$0.87 \pm 0.11 \pm 0.14$	<sup>2,3</sup> AAD	14BB ATLS	S-ratio in $\ell\ell + \geq 2j$ events
$0.75 \pm 0.19 \pm 0.23$	<sup>2,4</sup> AAD	14BB ATLS	$\cos\theta(\ell^+ \cos\theta(\ell^-))$ in $\ell\ell + \geq 2j$ events
$0.83 \pm 0.14 \pm 0.18$	<sup>2,5</sup> AAD	14BB ATLS	$\cos\theta(\ell^+ \cos\theta(\ell^-))$ in $\ell\ell + \geq 2j$ events

- <sup>1</sup> AAD 15J based on  $20.3 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 8 \text{ TeV}$ . Uses a fit including a linear superposition of  $\Delta\phi$  distribution from the SM NLO simulation with coefficient  $f_{SM}$  and from  $t\bar{t}$  simulation without spin correlation with coefficient  $(1 - f_{SM})$ .
- <sup>2</sup> Based on  $4.6 \text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 7 \text{ TeV}$ . The results are for  $m_t = 172.5 \text{ GeV}$ .
- <sup>3</sup> The S-ratio is defined as the SM spin correlation in the like-helicity gluon-gluon collisions normalized to the no spin correlation case; see eq.(6) for the LO expression.
- <sup>4</sup> The polar angle correlation along the helicity axis.
- <sup>5</sup> The polar angle correlation along the direction which maximizes the correlation.

**t-quark FCNC Couplings  $\kappa^{utg}/\Lambda$  and  $\kappa^{ctg}/\Lambda$** 

VALUE (TeV <sup>-1</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.0069$	95	<sup>1</sup> AAD	12BP ATLS	$t^{tug}/\Lambda$ ( $t^{tgc} = 0$ )
$< 0.016$	95	<sup>1</sup> AAD	12BP ATLS	$t^{tgc}/\Lambda$ ( $t^{tug} = 0$ )
$< 0.013$	95	<sup>2</sup> ABAZOV	10K D0	$\kappa^{tug}/\Lambda$
$< 0.057$	95	<sup>2</sup> ABAZOV	10K D0	$\kappa^{tgc}/\Lambda$
$< 0.018$	95	<sup>3</sup> AALTONEN	09N CDF	$\kappa^{tug}/\Lambda$ ( $\kappa^{tgc} = 0$ )
$< 0.069$	95	<sup>3</sup> AALTONEN	09N CDF	$\kappa^{tgc}/\Lambda$ ( $\kappa^{tug} = 0$ )
$< 0.037$	95	<sup>4</sup> ABAZOV	07V D0	$\kappa^{utg}/\Lambda$
$< 0.15$	95	<sup>4</sup> ABAZOV	07V D0	$\kappa^{ctg}/\Lambda$

# Quark Particle Listings

*t*

- <sup>1</sup> Based on 2.05 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. The results are obtained from the 95% CL upper limit on the single top-quark production  $\sigma(qg \rightarrow t)\text{-}B(t \rightarrow bW) < 3.9$  pb, for  $q=u$  or  $q=c$ ,  $B(t \rightarrow ug) < 5.7 \times 10^{-5}$  and  $B(t \rightarrow ug) < 2.7 \times 10^{-4}$ .
- <sup>2</sup> Based on 2.3 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. Upper limit of single top quark production cross section 0.20 pb and 0.27 pb via FCNC  $t$ - $u$ - $g$  and  $t$ - $c$ - $g$  couplings, respectively, lead to the bounds without assuming the absence of the other coupling.  $B(t \rightarrow u + g) < 2.0 \times 10^{-4}$  and  $B(t \rightarrow c + g) < 3.9 \times 10^{-3}$  follow.
- <sup>3</sup> Based on 2.2 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. Upper limit of single top quark production cross section  $\sigma(u(c) + g \rightarrow t) < 1.8$  pb (95% CL) via FCNC  $t$ - $u$ - $g$  and  $t$ - $c$ - $g$  couplings lead to the bounds.  $B(t \rightarrow u + g) < 3.9 \times 10^{-4}$  and  $B(t \rightarrow c + g) < 5.7 \times 10^{-3}$  follow.
- <sup>4</sup> Result is based on 230 pb<sup>-1</sup> of data at  $\sqrt{s} = 1.96$  TeV. Absence of single top quark production events via FCNC  $t$ - $u$ - $g$  and  $t$ - $c$ - $g$  couplings lead to the upper bounds on the dimensioned couplings,  $\kappa_t^{u\bar{t}g}/\Lambda$  and  $\kappa_t^{c\bar{t}g}/\Lambda$ , respectively.

## $\sigma(Ht\bar{t})/\sigma(Ht\bar{t})_{SM}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<6.7	95	<sup>1</sup> AAD	15 ATLS	$H\bar{t}\bar{t}; H \rightarrow \gamma\gamma$
2.8 ± 1.0		<sup>2</sup> KHACHATRY...14H	CMS	$H \rightarrow b\bar{b}, \tau\bar{\tau}, \tau\gamma, W W/Z Z$ (leptons)

- <sup>1</sup> Based on 4.5 fb<sup>-1</sup> of data at 7 TeV and 20.3 fb<sup>-1</sup> at 8 TeV. The result is for  $m_H = 125.4$  GeV. The measurement constrains the top quark Yukawa coupling strength parameter  $\kappa_t = Y_t/Y_t^{SM}$  to be  $-1.3 < \kappa_t < 8.0$  (95% CL).

- <sup>2</sup> Based on 5.1 fb<sup>-1</sup> of  $pp$  data at 7 TeV and 19.7 fb<sup>-1</sup> at 8 TeV. The results are obtained by assuming the SM decay branching fractions for the Higgs boson of mass 125.6 GeV. The signal strength for individual Higgs decay channels are given in Fig. 13, and the preferred region in the  $(\kappa_V, \kappa_f)$  space is given in Fig. 14.

## Single $t$ -Quark Production Cross Section in $p\bar{p}$ Collisions at $\sqrt{s} = 1.8$ TeV

Direct probe of the  $t\bar{b}W$  coupling and possible new physics at  $\sqrt{s} = 1.8$  TeV.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<24	95	<sup>1</sup> ACOSTA	04H CDF	$p\bar{p} \rightarrow t\bar{b} + X, tqb + X$
<18	95	<sup>2</sup> ACOSTA	02 CDF	$p\bar{p} \rightarrow t\bar{b} + X$
<13	95	<sup>3</sup> ACOSTA	02 CDF	$p\bar{p} \rightarrow tqb + X$

- <sup>1</sup> ACOSTA 04H bounds single top-quark production from the  $s$ -channel  $W$ -exchange process,  $q'\bar{q} \rightarrow t\bar{b}$ , and the  $t$ -channel  $W$ -exchange process,  $q'g \rightarrow q\bar{t}\bar{b}$ . Based on  $\sim 106$  pb<sup>-1</sup> of data.

- <sup>2</sup> ACOSTA 02 bounds the cross section for single top-quark production via the  $s$ -channel  $W$ -exchange process,  $q'\bar{q} \rightarrow t\bar{b}$ . Based on  $\sim 106$  pb<sup>-1</sup> of data.

- <sup>3</sup> ACOSTA 02 bounds the cross section for single top-quark production via the  $t$ -channel  $W$ -exchange process,  $q'g \rightarrow q\bar{t}\bar{b}$ . Based on  $\sim 106$  pb<sup>-1</sup> of data.

## Single $t$ -Quark Production Cross Section in $p\bar{p}$ Collisions at $\sqrt{s} = 1.96$ TeV

Direct probes of the  $t\bar{b}W$  coupling and possible new physics at  $\sqrt{s} = 1.96$  TeV.

OUR AVERAGE assumes that the systematic uncertainties are uncorrelated.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.25 <sup>+0.29</sup> <sub>-0.31</sub>		<sup>1</sup> AALTONEN	15H TEVA	$t$ -channel
3.30 <sup>+0.52</sup> <sub>-0.40</sub>		<sup>1,2</sup> AALTONEN	15H TEVA	$s$ - + $t$ -channels
1.12 <sup>+0.61</sup> <sub>-0.57</sub>		<sup>3</sup> AALTONEN	14K CDF	$s$ -channel ( $0\ell + \cancel{E}_T + 2, 3j$ ( $\geq 1b$ -tag))
1.41 <sup>+0.44</sup> <sub>-0.42</sub>		<sup>4</sup> AALTONEN	14L CDF	$s$ -channel ( $\ell + \cancel{E}_T + 2j$ ( $\geq 1b$ -tag))
1.29 <sup>+0.26</sup> <sub>-0.24</sub>		<sup>5</sup> AALTONEN	14M TEVA	$s$ -channel (CDF + D0)
3.04 <sup>+0.57</sup> <sub>-0.53</sub>		<sup>6</sup> AALTONEN	14o CDF	$s$ + $t$ + $Wt$ ( $\ell + \cancel{E}_T + 2$ or 3 jets ( $\geq 1b$ -tag))
1.10 <sup>+0.33</sup> <sub>-0.31</sub>		<sup>7</sup> ABAZOV	13o D0	$s$ -channel
3.07 <sup>+0.54</sup> <sub>-0.49</sub>		<sup>7</sup> ABAZOV	13o D0	$t$ -channel
4.11 <sup>+0.60</sup> <sub>-0.55</sub>		<sup>7</sup> ABAZOV	13o D0	$s$ - + $t$ -channels
0.98 ± 0.63		<sup>8</sup> ABAZOV	11AA D0	$s$ -channel
2.90 ± 0.59		<sup>8</sup> ABAZOV	11AA D0	$t$ -channel
3.43 <sup>+0.73</sup> <sub>-0.74</sub>		<sup>9</sup> ABAZOV	11AD D0	$s$ - + $t$ -channels
1.8 <sup>+0.7</sup> <sub>-0.5</sub>		<sup>10</sup> AALTONEN	10AB CDF	$s$ -channel
0.8 ± 0.4		<sup>10</sup> AALTONEN	10AB CDF	$t$ -channel
4.9 <sup>+2.5</sup> <sub>-2.2</sub>		<sup>11</sup> AALTONEN	10U CDF	$\cancel{E}_T$ + jets decay
3.14 <sup>+0.94</sup> <sub>-0.80</sub>		<sup>12</sup> ABAZOV	10 D0	$t$ -channel
1.05 ± 0.81		<sup>12</sup> ABAZOV	10 D0	$s$ -channel
< 7.3	95	<sup>13</sup> ABAZOV	10J D0	$\tau$ + jets decay
2.3 <sup>+0.6</sup> <sub>-0.5</sub>		<sup>14</sup> AALTONEN	09AT CDF	$s$ - + $t$ -channel
3.94 ± 0.88		<sup>15</sup> ABAZOV	09Z D0	$s$ - + $t$ -channel
2.2 <sup>+0.7</sup> <sub>-0.6</sub>		<sup>16</sup> AALTONEN	08AH CDF	$s$ - + $t$ -channel
4.7 ± 1.3		<sup>17</sup> ABAZOV	08i D0	$s$ - + $t$ -channel
4.9 ± 1.4		<sup>18</sup> ABAZOV	07H D0	$s$ - + $t$ -channel
< 6.4	95	<sup>19</sup> ABAZOV	05P D0	$p\bar{p} \rightarrow t\bar{b} + X$
< 5.0	95	<sup>19</sup> ABAZOV	05P D0	$p\bar{p} \rightarrow tqb + X$
< 10.1	95	<sup>20</sup> ACOSTA	05N CDF	$p\bar{p} \rightarrow tqb + X$
< 13.6	95	<sup>20</sup> ACOSTA	05N CDF	$p\bar{p} \rightarrow t\bar{b} + X$
< 17.8	95	<sup>20</sup> ACOSTA	05N CDF	$p\bar{p} \rightarrow t\bar{b} + X, tqb + X$

- <sup>1</sup> AALTONEN 15H based on 9.7 fb<sup>-1</sup> of data per experiment. The result is for  $m_t = 172.5$  GeV, and is a combination of the CDF measurements (AALTONEN 16) and the D0 measurements (ABAZOV 13o) on the  $t$ -channel single  $t$ -quark production cross section. The result is consistent with the NLO+NNLL SM prediction and gives  $|V_{tb}| = 1.02^{+0.06}_{-0.05}$  and  $|V_{tb}| > 0.92$  (95% CL).

- <sup>2</sup> AALTONEN 15H is a combined measurement of  $s$ -channel single top cross section by CDF + D0. AALTONEN 14M is not included.

- <sup>3</sup> Based on 9.45 fb<sup>-1</sup> of data, using neural networks to separate signal from backgrounds. The result is for  $m_t = 172.5$  GeV. Combination of this result with the CDF measurement in the 1 lepton channel AALTONEN 14L gives  $1.36^{+0.37}_{-0.32}$  pb, consistent with the SM prediction, and is 4.2 sigma away from the background only hypothesis.

- <sup>4</sup> Based on 9.4 fb<sup>-1</sup> of data, using neural networks to separate signal from backgrounds. The result is for  $m_t = 172.5$  GeV. The result is 3.8 sigma away from the background only hypothesis.

- <sup>5</sup> Based on 9.7 fb<sup>-1</sup> of data per experiment. The result is for  $m_t = 172.5$  GeV, and is a combination of the CDF measurements AALTONEN 14L, AALTONEN 14K and the D0 measurement ABAZOV 13o on the  $s$ -channel single  $t$ -quark production cross section. The result is consistent with the SM prediction of  $1.05 \pm 0.06$  pb and the significance of the observation is of 6.3 standard deviations.

- <sup>6</sup> Based on 7.5 fb<sup>-1</sup> of data. Neural network is used to discriminate signals ( $s$ -,  $t$ - and  $Wt$ -channel single top production) from backgrounds. The result is consistent with the SM prediction, and gives  $|V_{tb}| = 0.95 \pm 0.09$  (stat + syst)  $\pm 0.05$  (theory) and  $|V_{tb}| > 0.78$  (95% CL). The result is for  $m_t = 172.5$  GeV.

- <sup>7</sup> Based on 9.7 fb<sup>-1</sup> of data. Events with  $\ell + \cancel{E}_T + 2$  or 3 jets (1 or 2  $b$ -tag) are analysed, assuming  $m_t = 172.5$  GeV. The combined  $s$ - +  $t$ -channel cross section gives  $|V_{tb}| \frac{f_L^1}{f_L^1} = 1.12^{+0.09}_{-0.08}$ , or  $|V_{tb}| > 0.92$  at 95% CL for  $f_L^1 = 1$  and a flat prior within  $0 \leq |V_{tb}|^2 \leq 1$ .

- <sup>8</sup> Based on 5.4 fb<sup>-1</sup> of data. The error is statistical + systematic combined. The results are for  $m_t = 172.5$  GeV. Results for other  $m_t$  values are given in Table 2 of ABAZOV 11AA.

- <sup>9</sup> Based on 5.4 fb<sup>-1</sup> of data and for  $m_t = 172.5$  GeV. The error is statistical + systematic combined. Results for other  $m_t$  values are given in Table III of ABAZOV 11AD. The result is obtained by assuming the SM ratio between  $t\bar{b}$  ( $s$ -channel) and  $tqb$  ( $t$ -channel) productions, and gives  $|V_{tb}| \frac{f_L^1}{f_L^1} = 1.02^{+0.10}_{-0.11}$ , or  $|V_{tb}| > 0.79$  at 95% CL for a flat prior within  $0 < |V_{tb}|^2 < 1$ .

- <sup>10</sup> Based on 3.2 fb<sup>-1</sup> of data. For combined  $s$ - +  $t$ -channel result see AALTONEN 09AT.

- <sup>11</sup> Result is based on 2.1 fb<sup>-1</sup> of data. Events with large missing  $E_T$  and jets with at least one  $b$ -jet without identified electron or muon are selected. Result is obtained when observed  $2.1 \sigma$  excess over the background originates from the signal for  $m_t = 175$  GeV, giving  $|V_{tb}| = 1.24^{+0.34}_{-0.29} \pm 0.07$  (theory).

- <sup>12</sup> Result is based on 2.3 fb<sup>-1</sup> of data. Events with isolated  $\ell + \cancel{E}_T + 2, 3, 4$  jets with one or two  $b$ -tags are selected. The analysis assumes  $m_t = 170$  GeV.

- <sup>13</sup> Result is based on 4.8 fb<sup>-1</sup> of data. Events with an isolated reconstructed tau lepton, missing  $E_T + 2, 3$  jets with one or two  $b$ -tags are selected. When combined with ABAZOV 09Z result for  $e + \mu$  channels, the  $s$ - and  $t$ -channels combined cross section is  $3.84^{+0.69}_{-0.83}$  pb.

- <sup>14</sup> Based on 3.2 fb<sup>-1</sup> of data. Events with isolated  $\ell + \cancel{E}_T$  + jets with at least one  $b$ -tag are analyzed and  $s$ - and  $t$ -channel single top events are selected by using the likelihood function, matrix element, neural-network, boosted decision tree, likelihood function optimized for  $s$ -channel process, and neural-networked based analysis of events with  $\cancel{E}_T$  that has sensitivity for  $W \rightarrow \tau\nu$  decays. The result is for  $m_t = 175$  GeV, and the mean value decreases by 0.02 pb/GeV for smaller  $m_t$ . The signal has 5.0 sigma significance. The result gives  $|V_{tb}| = 0.91 \pm 0.11$  (stat+syst)  $\pm 0.07$  (theory), or  $|V_{tb}| > 0.71$  at 95% CL.

- <sup>15</sup> Based on 2.3 fb<sup>-1</sup> of data. Events with isolated  $\ell + \cancel{E}_T + \geq 2$  jets with 1 or 2  $b$ -tags are analyzed and  $s$ - and  $t$ -channel single top events are selected by using boosted decision tree, Bayesian neural networks and the matrix element method. The signal has 5.0 sigma significance. The result gives  $|V_{tb}| = 1.07 \pm 0.12$ , or  $|V_{tb}| > 0.78$  at 95% CL. The analysis assumes  $m_t = 170$  GeV.

- <sup>16</sup> Result is based on 2.2 fb<sup>-1</sup> of data. Events with isolated  $\ell + \cancel{E}_T + 2, 3$  jets with at least one  $b$ -tag are selected, and  $s$ - and  $t$ -channel single top events are selected by using likelihood, matrix element, and neural network discriminants. The result can be interpreted as  $|V_{tb}| = 0.88^{+0.13}_{-0.12}$  (stat + syst)  $\pm 0.07$  (theory), and  $|V_{tb}| > 0.66$  (95% CL) under the  $|V_{tb}| < 1$  constraint.

- <sup>17</sup> Result is based on 0.9 fb<sup>-1</sup> of data. Events with isolated  $\ell + \cancel{E}_T + 2, 3, 4$  jets with one or two  $b$ -vertex-tag are selected, and contributions from  $W$  + jets,  $t\bar{t}$ ,  $s$ - and  $t$ -channel single top events are identified by using boosted decision trees, Bayesian neural networks, and matrix element analysis. The result can be interpreted as the measurement of the CKM matrix element  $|V_{tb}| = 1.31^{+0.25}_{-0.21}$ , or  $|V_{tb}| > 0.68$  (95% CL) under the  $|V_{tb}| < 1$  constraint.

- <sup>18</sup> Result is based on 0.9 fb<sup>-1</sup> of data. This result constrains  $V_{tb}$  to  $0.68 < |V_{tb}| \leq 1$  at 95% CL.

- <sup>19</sup> ABAZOV 05P bounds single top-quark production from either the  $s$ -channel  $W$ -exchange process,  $q'\bar{q} \rightarrow t\bar{b}$ , or the  $t$ -channel  $W$ -exchange process,  $q'g \rightarrow q\bar{t}\bar{b}$ , based on  $\sim 230$  pb<sup>-1</sup> of data.

- <sup>20</sup> ACOSTA 05N bounds single top-quark production from the  $t$ -channel  $W$ -exchange process ( $q'g \rightarrow q\bar{t}\bar{b}$ ), the  $s$ -channel  $W$ -exchange process ( $q'\bar{q} \rightarrow t\bar{b}$ ), and from the combined cross section of  $t$ - and  $s$ -channel. Based on  $\sim 162$  pb<sup>-1</sup> of data.

## $t$ -channel Single $t$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 7$ TeV

Direct probe of the  $t\bar{b}W$  coupling and possible new physics at  $\sqrt{s} = 7$  TeV.

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
68 ± 2 ± 8	<sup>1</sup> AAD	14B1 ATLS	$\ell + \cancel{E}_T + 2j$ or $3j$
83 ± 4 ± 20	<sup>2</sup> AAD	12CH ATLS	$t$ -channel $\ell + \cancel{E}_T + (2,3)j$ (1b)
67.2 ± 6.1	<sup>3</sup> CHATRCHYAN12BQ	CMS	$t$ -channel $\ell + \cancel{E}_T + \geq 2j$ (1b)
83.6 ± 29.8 ± 3.3	<sup>4</sup> CHATRCHYAN11R	CMS	$t$ -channel

<sup>1</sup> Based on  $4.59 \text{ fb}^{-1}$  of data, using neural networks for signal and background separation.  $\sigma(tq) = 46 \pm 1 \pm 6 \text{ pb}$  and  $\sigma(\bar{t}q) = 23 \pm 1 \pm 3 \text{ pb}$  are separately measured, as well as their ratio  $R = \sigma(tq)/\sigma(\bar{t}q) = 2.04 \pm 0.13 \pm 0.12$ . The results are for  $m_t = 172.5 \text{ GeV}$ , and those for other  $m_t$  values are given by eq.(4) and Table IV. The measurements give  $|V_{tb}| = 1.02 \pm 0.07$  or  $|V_{tb}| > 0.88$  (95% CL).

<sup>2</sup> Based on  $1.04 \text{ fb}^{-1}$  of data. The result gives  $|V_{tb}| = 1.13^{+0.14}_{-0.13}$  from the ratio  $\sigma(\text{exp})/\sigma(\text{th})$ , where  $\sigma(\text{th})$  is the SM prediction for  $|V_{tb}| = 1$ . The 95% CL lower bound of  $|V_{tb}| > 0.75$  is found if  $|V_{tb}| < 1$  is assumed.  $\sigma(t) = 59^{+18}_{-16} \text{ pb}$  and  $\sigma(\bar{t}) = 33^{+13}_{-12} \text{ pb}$  are found for the separate single  $t$  and  $\bar{t}$  production cross sections, respectively. The results assume  $m_t = 172.5 \text{ GeV}$  for the acceptance.

<sup>3</sup> Based on  $1.17 \text{ fb}^{-1}$  of data for  $\ell = \mu$ ,  $1.56 \text{ fb}^{-1}$  of data for  $\ell = e$  at  $7 \text{ TeV}$  collected during 2011. The result gives  $|V_{tb}| = 1.020 \pm 0.046(\text{meas}) \pm 0.017(\text{th})$ . The 95% CL lower bound of  $|V_{tb}| > 0.92$  is found if  $|V_{tb}| < 1$  is assumed. The results assume  $m_t = 172.5 \text{ GeV}$  for the acceptance.

<sup>4</sup> Based on  $36 \text{ pb}^{-1}$  of data. The first error is statistical + systematic combined, the second is luminosity. The result gives  $|V_{tb}| = 1.114 \pm 0.22(\text{exp}) \pm 0.02(\text{th})$  from the ratio  $\sigma(\text{exp})/\sigma(\text{th})$ , where  $\sigma(\text{th})$  is the SM prediction for  $|V_{tb}| = 1$ . The 95% CL lower bound of  $|V_{tb}| > 0.62$  (0.68) is found from the 2D (BDT) analysis under the constraint  $0 < |V_{tb}|^2 < 1$ .

### W t Production Cross Section in pp Collisions at $\sqrt{s} = 7 \text{ TeV}$

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$16^{+5}_{-4}$	<sup>1</sup> CHATRCHYAN13C	CMS	$t + W$ channel, $2\ell + \cancel{E}_T + 1b$
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<sup>1</sup> Based on  $4.9 \text{ fb}^{-1}$  of data. The result gives  $V_{tb} = 1.01^{+0.16}_{-0.13}(\text{exp})^{+0.03}_{-0.04}(\text{th})$ .  $V_{tb} > 0.79$  (95% CL) if  $V_{tb} < 1$  is assumed. The results assume  $m_t = 172.5 \text{ GeV}$  for the acceptance.

### t-channel Single t Production Cross Section in pp Collisions at $\sqrt{s} = 8 \text{ TeV}$

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$83.6 \pm 2.3 \pm 7.4$	<sup>1</sup> KHACHATRY...14F	CMS	$\ell + \cancel{E}_T + \geq 2j$ (1,2 b, 1 forward j)
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<sup>1</sup> Based on  $19.7 \text{ fb}^{-1}$  of data. The  $t$  and  $\bar{t}$  production cross sections are measured separately as  $\sigma_{t\text{-ch.}}(t) = 53.8 \pm 1.5 \pm 4.4 \text{ pb}$  and  $\sigma_{t\text{-ch.}}(\bar{t}) = 27.6 \pm 1.3 \pm 3.7 \text{ pb}$ , respectively, as well as their ratio  $R_{t\text{-ch.}} = \sigma_{t\text{-ch.}}(t)/\sigma_{t\text{-ch.}}(\bar{t}) = 1.95 \pm 0.10 \pm 0.19$ , in agreement with the SM predictions. Combination with a previous CMS result at  $\sqrt{s} = 7 \text{ TeV}$  [CHATRCHYAN 12BQ] gives  $|V_{tb}| = 0.998 \pm 0.038 \pm 0.016$ . Also obtained is the ratio  $R_{8/7} = \sigma_{t\text{-ch.}}(8\text{TeV})/\sigma_{t\text{-ch.}}(7\text{TeV}) = 1.24 \pm 0.08 \pm 0.12$ .

### s-channel Single t Production Cross Section in pp Collisions at $\sqrt{s} = 8 \text{ TeV}$

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$5.0 \pm 4.3$	<sup>1</sup> AAD	15A ATLS	$\ell + \cancel{E}_T + 2b$
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<sup>1</sup> Based on  $20.3 \text{ fb}^{-1}$  of data, using a multivariate analysis to separate signal and backgrounds. The 95% CL upper bound of the cross section is  $14.6 \text{ pb}$ . The results are consistent with the SM prediction of  $5.61 \pm 0.22 \text{ pb}$  at approximate NNLO.

### W t Production Cross Section in pp Collisions at $\sqrt{s} = 8 \text{ TeV}$

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$23.0 \pm 1.3^{+3.2}_{-3.5} \pm 1.1$	<sup>1</sup> AAD	16B ATLS	$2\ell + \cancel{E}_T + 1b$
$23.4 \pm 5.4$	<sup>2</sup> CHATRCHYAN14AC	CMS	$t + W$ channel, $2\ell + \cancel{E}_T + 1b$

<sup>1</sup> AAD 16B based on  $20.3 \text{ fb}^{-1}$  of data. The result gives  $|V_{tb}| = 1.01 \pm 0.10$  and  $|V_{tb}| > 0.80$  (95% CL) without assuming unitarity of the CKM matrix. The results assume  $m_t = 172.5 \text{ GeV}$  for the acceptance.

<sup>2</sup> Based on  $12.2 \text{ fb}^{-1}$  of data. Events with two oppositely charged leptons, large  $\cancel{E}_T$  and a  $b$ -tagged jet are selected, and a multivariate analysis is used to separate the signal from the backgrounds. The result is consistent with the SM prediction of  $22.2 \pm 0.6(\text{scale}) \pm 1.4(\text{PDF}) \text{ pb}$  at approximate NNLO.

### Single t-Quark Production Cross Section in ep Collisions

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 0.25$	95	<sup>1</sup> AARON	09A H1	$e^\pm p \rightarrow e^\pm t X$
$< 0.55$	95	<sup>2</sup> AKTAS	04 H1	$e^\pm p \rightarrow e^\pm t X$
$< 0.225$	95	<sup>3</sup> CHEKANOV	03 ZEUS	$e^\pm p \rightarrow e^\pm t X$

<sup>1</sup> AARON 09A looked for single top production via FCNC in  $e^\pm p$  collisions at HERA with  $474 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 301\text{--}319 \text{ GeV}$ . The result supersedes that of AKTAS 04.

<sup>2</sup> AKTAS 04 looked for single top production via FCNC in  $e^\pm$  collisions at HERA with  $118.3 \text{ pb}^{-1}$ , and found 5 events in the  $e$  or  $\mu$  channels while  $1.31 \pm 0.22$  events are expected from the Standard Model background. No excess was found for the hadronic channel. The observed cross section of  $\sigma(ep \rightarrow etX) = 0.29^{+0.15}_{-0.14} \text{ pb}$  at  $\sqrt{s} = 319 \text{ GeV}$  gives the quoted upper bound if the observed events are due to statistical fluctuation.

<sup>3</sup> CHEKANOV 03 looked in  $130.1 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 301$  and  $318 \text{ GeV}$ . The limit is for  $\sqrt{s} = 318 \text{ GeV}$  and assumes  $m_t = 175 \text{ GeV}$ .

### t $\bar{t}$ Production Cross Section in p $\bar{p}$ Collisions at $\sqrt{s} = 1.8 \text{ TeV}$

Only the final combined  $t\bar{t}$  production cross sections obtained from Tevatron Run I by the CDF and D0 experiments are quoted below.

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$5.69 \pm 1.21 \pm 1.04$	<sup>1</sup> ABAZOV	03A D0	Combined Run I data
$6.5^{+1.7}_{-1.4}$	<sup>2</sup> AFFOLDER	01A CDF	Combined Run I data

<sup>1</sup> Combined result from  $110 \text{ pb}^{-1}$  of Tevatron Run I data. Assume  $m_t = 172.1 \text{ GeV}$ .

<sup>2</sup> Combined result from  $105 \text{ pb}^{-1}$  of Tevatron Run I data. Assume  $m_t = 175 \text{ GeV}$ .

### t $\bar{t}$ Production Cross Section in p $\bar{p}$ Collisions at $\sqrt{s} = 1.96 \text{ TeV}$

Unless otherwise noted the first quoted error is from statistics, the second from systematic uncertainties, and the third from luminosity. If only two errors are quoted the luminosity is included in the systematic uncertainties.

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$8.1 \pm 2.1$	<sup>1</sup> AALTONEN	14A CDF	$\ell + \cancel{E}_T + \geq 2\text{jets}$ ( $\geq 1b$ -tag)
$7.60 \pm 0.20 \pm 0.29 \pm 0.21$	<sup>2</sup> AALTONEN	14H TEVA	$\ell\ell, \ell + \text{jets}$ , all-jets channels
$8.0 \pm 0.7 \pm 0.6 \pm 0.5$	<sup>3</sup> ABAZOV	14K D0	$\ell + \cancel{E}_T + \geq 4\text{jets}$ ( $\geq 1b$ -tag)
$7.09 \pm 0.84$	<sup>4</sup> AALTONEN	13AB CDF	$\ell\ell + \cancel{E}_T + \geq 2\text{jets}$
$7.5 \pm 1.0$	<sup>5</sup> AALTONEN	13G CDF	$\ell + \cancel{E}_T + \geq 3\text{jets}$ ( $\geq 1b$ -tag)
$8.8 \pm 3.3 \pm 2.2$	<sup>6</sup> AALTONEN	12AL CDF	$\cancel{E}_T + \cancel{E}_T + 4j$ ( $\geq 1b$ )
$8.5 \pm 0.6 \pm 0.7$	<sup>7</sup> AALTONEN	11D CDF	$\ell + \cancel{E}_T + \text{jets}$ ( $\geq 1b$ -tag)
$7.64 \pm 0.57 \pm 0.45$	<sup>8</sup> AALTONEN	11W CDF	$\ell + \cancel{E}_T + \text{jets}$ ( $\geq 1b$ -tag)
$7.99 \pm 0.55 \pm 0.76 \pm 0.46$	<sup>9</sup> AALTONEN	11Y CDF	$\cancel{E}_T + \geq 4\text{jets}$ (0,1,2 b-tag)
$7.78^{+0.77}_{-0.64}$	<sup>10</sup> ABAZOV	11E D0	$\ell + \cancel{E}_T + \geq 2\text{jets}$
$7.56^{+0.63}_{-0.56}$	<sup>11</sup> ABAZOV	11Z D0	Combination
$6.27 \pm 0.73 \pm 0.63 \pm 0.39$	<sup>12</sup> AALTONEN	10AA CDF	Repl. by AALTONEN 13AB
$7.2 \pm 0.5 \pm 1.0 \pm 0.4$	<sup>13</sup> AALTONEN	10E CDF	$\geq 6\text{jets}$ , vtx b-tag
$7.8 \pm 2.4 \pm 1.6 \pm 0.5$	<sup>14</sup> AALTONEN	10V CDF	$\ell + \geq 3\text{jets}$ , soft- $e$ b-tag
$7.70 \pm 0.52$	<sup>15</sup> AALTONEN	10W CDF	$\ell + \cancel{E}_T + \geq 3\text{jets} + b$ -tag, norm. to $\sigma(Z \rightarrow \ell\ell)_{TH}$
$6.9 \pm 2.0$	<sup>16</sup> ABAZOV	10I D0	$\geq 6\text{jets}$ with 2 b-tags
$6.9 \pm 1.2^{+0.8}_{-0.7} \pm 0.4$	<sup>17</sup> ABAZOV	10Q D0	$\tau_h + \text{jets}$
$9.6 \pm 1.2^{+0.6}_{-0.5} \pm 0.6$	<sup>18</sup> AALTONEN	09AD CDF	$\ell\ell + \cancel{E}_T$ / vtx b-tag
$9.1 \pm 1.1^{+1.0}_{-0.9} \pm 0.6$	<sup>19</sup> AALTONEN	09H CDF	$\ell + \geq 3\text{jets} + \cancel{E}_T$ / soft $\mu$ b-tag
$8.18^{+0.98}_{-0.87}$	<sup>20</sup> ABAZOV	09AG D0	$\ell + \text{jets}$ , $\ell\ell$ and $\ell\tau + \text{jets}$
$7.5 \pm 1.0^{+0.7}_{-0.6} \pm 0.5$	<sup>21</sup> ABAZOV	09R D0	$\ell\ell$ and $\ell\tau + \text{jets}$
$8.18^{+0.90}_{-0.84} \pm 0.50$	<sup>22</sup> ABAZOV	08M D0	$\ell + n\text{jets}$ with 0,1,2 b-tag
$7.62 \pm 0.85$	<sup>23</sup> ABAZOV	08N D0	$\ell + n\text{jets} + b$ -tag or kinematics
$8.5^{+2.7}_{-2.2}$	<sup>24</sup> ABULENCIA	08 CDF	$\ell^+ \ell^-$ ( $\ell = e, \mu$ )
$8.3 \pm 1.0^{+2.0}_{-1.5} \pm 0.5$	<sup>25</sup> AALTONEN	07D CDF	$\geq 6\text{jets}$ , vtx b-tag
$7.4 \pm 1.4 \pm 1.0$	<sup>26</sup> ABAZOV	07O D0	$\ell\ell + \text{jets}$ , vtx b-tag
$4.5^{+2.0}_{-1.9} \pm 1.4 \pm 0.3$	<sup>27</sup> ABAZOV	07P D0	$\geq 6\text{jets}$ , vtx b-tag
$6.4^{+1.3}_{-1.2} \pm 0.7 \pm 0.4$	<sup>28</sup> ABAZOV	07R D0	$\ell + \geq 4\text{jets}$
$6.6 \pm 0.9 \pm 0.4$	<sup>29</sup> ABAZOV	06X D0	$\ell + \text{jets}$ , vtx b-tag
$8.7 \pm 0.9^{+1.1}_{-0.9}$	<sup>30</sup> ABULENCIA	06Z CDF	$\ell + \text{jets}$ , vtx b-tag
$5.8 \pm 1.2^{+0.9}_{-0.7}$	<sup>31</sup> ABULENCIA,A	06C CDF	missing $E_T + \text{jets}$ , vtx b-tag
$7.5 \pm 2.1^{+3.3}_{-2.2} \pm 0.5$	<sup>32</sup> ABULENCIA,A	06E CDF	6–8 jets, b-tag
$8.9 \pm 1.0^{+1.1}_{-1.0}$	<sup>33</sup> ABULENCIA,A	06F CDF	$\ell + \geq 3\text{jets}$ , b-tag
$8.6^{+1.6}_{-1.5} \pm 0.6$	<sup>34</sup> ABAZOV	05Q D0	$\ell + n\text{jets}$
$8.6^{+3.2}_{-2.7} \pm 1.1 \pm 0.6$	<sup>35</sup> ABAZOV	05R D0	di-lepton + n jets
$6.7^{+1.4}_{-1.3} \pm 1.6 \pm 0.4$	<sup>36</sup> ABAZOV	05X D0	$\ell + \text{jets}$ / kinematics
$5.3 \pm 3.3^{+1.3}_{-1.0}$	<sup>37</sup> ACOSTA	05S CDF	$\ell + \text{jets}$ / soft $\mu$ b-tag
$6.6 \pm 1.1 \pm 1.5$	<sup>38</sup> ACOSTA	05T CDF	$\ell + \text{jets}$ / kinematics
$6.0^{+1.5}_{-1.6} \pm 1.2 \pm 1.3$	<sup>39</sup> ACOSTA	05U CDF	$\ell + \text{jets}$ / kinematics + vtx b-tag
$5.6^{+1.2}_{-1.1} \pm 0.9 \pm 0.6$	<sup>40</sup> ACOSTA	05V CDF	$\ell + n\text{jets}$
$7.0^{+2.4}_{-2.1} \pm 1.6 \pm 0.4$	<sup>41</sup> ACOSTA	04I CDF	di-lepton + jets + missing ET

<sup>1</sup> Based on  $9 \text{ fb}^{-1}$  of data. The measurement is in the channel  $t\bar{t} \rightarrow (b\ell\nu)(b\tau\nu)$ , where  $\tau$  decays into hadrons ( $\tau_h$ ), and  $\ell$  ( $e$  or  $\mu$ ) include  $\ell$  from  $\tau$  decays ( $\tau_\ell$ ). The result is for  $m_t = 173 \text{ GeV}$ .

<sup>2</sup> Based on  $8.8 \text{ fb}^{-1}$  of data. Combination of CDF and D0 measurements given, respectively, by  $\sigma(t\bar{t}; \text{CDF}) = 7.63 \pm 0.31 \pm 0.36 \pm 0.16 \text{ pb}$ ,  $\sigma(t\bar{t}; \text{D0}) = 7.56 \pm 0.20 \pm 0.32 \pm 0.46 \text{ pb}$ . All the results are for  $m_t = 172.5 \text{ GeV}$ . The  $m_t$  dependence of the mean value is parametrized in eq. (1) and shown in Fig. 2.

<sup>3</sup> Based on  $9.7 \text{ fb}^{-1}$  of data. Differential cross sections with respect to  $m_{t\bar{t}}$ ,  $|y(\text{top})|$ ,  $E_T(\text{top})$  are shown in Figs. 9, 10, 11, respectively, and are compared to the predictions of MC models.

<sup>4</sup> Based on  $8.8 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ .

<sup>5</sup> Based on  $8.7 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . Measure the  $t\bar{t}$  cross section simultaneously with the fraction of  $t \rightarrow Wb$  decays. The correlation coefficient between

# Quark Particle Listings

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- those two measurements is  $-0.434$ . Assume unitarity of the  $3 \times 3$  CKM matrix and set  $|V_{tb}| > 0.89$  at 95% CL.
- <sup>6</sup> Based on  $2.2 \text{ fb}^{-1}$  of data in  $p\bar{p}$  collisions at 1.96 TeV. The result assumes the acceptance for  $m_t = 172.5 \text{ GeV}$ .
- <sup>7</sup> Based on  $1.12 \text{ fb}^{-1}$  and assumes  $m_t = 175 \text{ GeV}$ , where the cross section changes by  $\pm 0.1 \text{ pb}$  for every  $\mp 1 \text{ GeV}$  shift in  $m_t$ . AALTONEN 11D fits simultaneously a jet production cross section and the  $b$ -tagging efficiency and find improvements in both measurements.
- <sup>8</sup> Based on  $2.7 \text{ fb}^{-1}$ . The first error is from statistics and systematics, the second is from luminosity. The result is for  $m_t = 175 \text{ GeV}$ . AALTONEN 11W fits simultaneously a jet flavor discriminator between  $b$ -,  $c$ -, and light-quarks, and find significant reduction in the systematic error.
- <sup>9</sup> Based on  $2.2 \text{ fb}^{-1}$ . The result is for  $m_t = 172.5 \text{ GeV}$ . AALTONEN 11Y selects multi-jet events with large  $E_T$ , and vetoes identified electrons and muons.
- <sup>10</sup> Based on  $5.3 \text{ fb}^{-1}$ . The error is statistical + systematic + luminosity combined. The result is for  $m_t = 172.5 \text{ GeV}$ . The results for other  $m_t$  values are given in Table XII and eq.(10) of ABAZOV 11E.
- <sup>11</sup> Combination of a dilepton measurement presented in ABAZOV 11Z (based on  $5.4 \text{ fb}^{-1}$ ), which yields  $7.36^{+0.90}_{-0.79} \text{ (stat+syst) pb}$ , and the lepton + jets measurement of ABAZOV 11E. The result is for  $m_t = 172.5 \text{ GeV}$ . The results for other  $m_t$  values is given by eq.(5) of ABAZOV 11A.
- <sup>12</sup> Based on  $2.8 \text{ fb}^{-1}$ . The result is for  $m_t = 175 \text{ GeV}$ .
- <sup>13</sup> Based on  $2.9 \text{ fb}^{-1}$ . Result is obtained from the fraction of signal events in the top quark mass measurement in the all hadronic decay channel.
- <sup>14</sup> Based on  $1.7 \text{ fb}^{-1}$ . The result is for  $m_t = 175 \text{ GeV}$ . AALTONEN 10V uses soft electrons from  $b$ -hadron decays to suppress  $W$ +jets background events.
- <sup>15</sup> Based on  $4.6 \text{ fb}^{-1}$ . The result is for  $m_t = 172.5 \text{ GeV}$ . The ratio  $\sigma(t\bar{t} \rightarrow \ell\text{+jets}) / \sigma(Z/\gamma^* \rightarrow \ell\bar{\ell})$  is measured and then multiplied by the theoretical  $Z/\gamma^* \rightarrow \ell\bar{\ell}$  cross section of  $\sigma(Z/\gamma^* \rightarrow \ell\bar{\ell}) = 251.3 \pm 5.0 \text{ pb}$ , which is free from the luminosity error.
- <sup>16</sup> Based on  $1 \text{ fb}^{-1}$ . The result is for  $m_t = 175 \text{ GeV}$ .  $7.9 \pm 2.3 \text{ pb}$  is found for  $m_t = 170 \text{ GeV}$ . ABAZOV 10I uses a likelihood discriminant to separate signal from background, where the background model was created from lower jet-multiplicity data.
- <sup>17</sup> Based on  $1 \text{ fb}^{-1}$ . The result is for  $m_t = 170 \text{ GeV}$ . For  $m_t = 175 \text{ GeV}$ , the result is  $6.3^{+1.2}_{-1.1} \text{ (stat)} \pm 0.7 \text{ (syst)} \pm 0.4 \text{ (lumi) pb}$ . Cross section of  $t\bar{t}$  production has been measured in the  $t\bar{t} \rightarrow \tau_h + \text{jets}$  topology, where  $\tau_h$  denotes hadronically decaying  $\tau$  leptons. The result for the cross section times the branching ratio is  $\sigma(t\bar{t}) \cdot B(t\bar{t} \rightarrow \tau_h + \text{jets}) = 0.60^{+0.23+0.15}_{-0.22-0.14} \pm 0.04 \text{ pb}$  for  $m_t = 170 \text{ GeV}$ .
- <sup>18</sup> Based on  $1.1 \text{ fb}^{-1}$ . The result is for  $B(W \rightarrow \ell\nu) = 10.8\%$  and  $m_t = 175 \text{ GeV}$ ; the mean value is  $9.8$  for  $m_t = 172.5 \text{ GeV}$  and  $10.1$  for  $m_t = 170 \text{ GeV}$ . AALTONEN 09AD used high  $p_T$   $e$  or  $\mu$  with an isolated track to select  $t\bar{t}$  decays into dileptons including  $\ell = \tau$ . The result is based on the candidate event samples with and without vertex  $b$ -tag.
- <sup>19</sup> Based on  $2 \text{ fb}^{-1}$ . The result is for  $m_t = 175 \text{ GeV}$ ; the mean value is 3% higher for  $m_t = 170 \text{ GeV}$  and 4% lower for  $m_t = 180 \text{ GeV}$ .
- <sup>20</sup> Result is based on  $1 \text{ fb}^{-1}$  of data. The result is for  $m_t = 170 \text{ GeV}$ , and the mean value decreases with increasing  $m_t$ ; see their Fig. 2. The result is obtained after combining  $\ell + \text{jets}$ ,  $\ell\bar{\ell}$ , and  $\ell\tau$  final states, and the ratios of the extracted cross sections are  $R^{\ell\bar{\ell}}/\ell j = 0.86^{+0.19}_{-0.17}$  and  $R^{\ell\tau}/\ell\bar{\ell} - \ell j = 0.97^{+0.32}_{-0.29}$ , consistent with the SM expectation of  $R = 1$ . This leads to the upper bound of  $B(t \rightarrow bH^+) \text{ as a function of } m_{H^+}$ . Results are shown in their Fig. 1 for  $B(H^+ \rightarrow \tau\nu) = 1$  and  $B(H^+ \rightarrow c\bar{s}) = 1$  cases. Comparison of the  $m_t$  dependence of the extracted cross section and a partial NNLO prediction gives  $m_t = 169.1^{+5.9}_{-5.2} \text{ GeV}$ .
- <sup>21</sup> Result is based on  $1 \text{ fb}^{-1}$  of data. The result is for  $m_t = 170 \text{ GeV}$ , and the mean value changes by  $-0.07 [m_t(\text{GeV}) - 170] \text{ pb}$  near the reference  $m_t$  value. Comparison of the  $m_t$  dependence of the extracted cross section and a partial NNLO QCD prediction gives  $m_t = 171.5^{+9.9}_{-8.8} \text{ GeV}$ . The  $\ell\tau$  channel alone gives  $7.6^{+4.9+3.5+1.4}_{-4.3-3.4-0.9} \text{ pb}$  and the  $\ell\bar{\ell}$  channel gives  $7.5^{+1.2+0.7+0.7}_{-1.1-0.6-0.5} \text{ pb}$ .
- <sup>22</sup> Result is based on  $0.9 \text{ fb}^{-1}$  of data. The first error is from stat + syst, while the latter error is from luminosity. The result is for  $m_t = 175 \text{ GeV}$ , and the mean value changes by  $-0.09 \text{ pb} [m_t(\text{GeV}) - 175]$ .
- <sup>23</sup> Result is based on  $0.9 \text{ fb}^{-1}$  of data. The cross section is obtained from the  $\ell + \geq 3$  jet event rates with 1 or 2  $b$ -tag, and also from the kinematical likelihood analysis of the  $\ell + 3, 4$  jet events. The result is for  $m_t = 172.6 \text{ GeV}$ , and its  $m_t$  dependence shown in Fig. 3 leads to the constraint  $m_t = 170 \pm 7 \text{ GeV}$  when compared to the SM prediction.
- <sup>24</sup> Result is based on  $360 \text{ pb}^{-1}$  of data. Events with high  $p_T$  oppositely charged dileptons  $\ell^+\ell^-$  ( $\ell = e, \mu$ ) are used to obtain cross sections for  $t\bar{t}$ ,  $W^+W^-$ , and  $Z \rightarrow \tau^+\tau^-$  production processes simultaneously. The other cross sections are given in Table IV.
- <sup>25</sup> Based on  $1.02 \text{ fb}^{-1}$  of data. Result is for  $m_t = 175 \text{ GeV}$ . Secondary vertex  $b$ -tag and neural network selections are used to achieve a signal-to-background ratio of about 1/2.
- <sup>26</sup> Based on  $425 \text{ pb}^{-1}$  of data. Result is for  $m_t = 175 \text{ GeV}$ . For  $m_t = 170.9 \text{ GeV}$ ,  $7.8 \pm 1.8 \text{ (stat + syst) pb}$  is obtained.
- <sup>27</sup> Based on  $405 \pm 25 \text{ pb}^{-1}$  of data. Result is for  $m_t = 175 \text{ GeV}$ . The last error is for luminosity. Secondary vertex  $b$ -tag and neural network are used to separate the signal events from the background.
- <sup>28</sup> Based on  $425 \text{ pb}^{-1}$  of data. Assumes  $m_t = 175 \text{ GeV}$ .
- <sup>29</sup> Based on  $\sim 425 \text{ pb}^{-1}$ . Assuming  $m_t = 175 \text{ GeV}$ . The first error is combined statistical and systematic, the second one is luminosity.
- <sup>30</sup> Based on  $\sim 318 \text{ pb}^{-1}$ . Assuming  $m_t = 178 \text{ GeV}$ . The cross section changes by  $\pm 0.08 \text{ pb}$  for each  $\mp 1 \text{ GeV}$  change in the assumed  $m_t$ . Result is for at least one  $b$ -tag. For at least two  $b$ -tagged jets,  $t\bar{t}$  signal of significance greater than  $5\sigma$  is found, and the cross section is  $10.1^{+1.6+2.0}_{-1.4-1.3} \text{ pb}$  for  $m_t = 178 \text{ GeV}$ .
- <sup>31</sup> Based on  $\sim 311 \text{ pb}^{-1}$ . Assuming  $m_t = 178 \text{ GeV}$ . For  $m_t = 175 \text{ GeV}$ , the result is  $6.0 \pm 1.2^{+0.9}_{-0.7}$ . This is the first CDF measurement without lepton identification, and hence it has sensitivity to the  $W \rightarrow \tau\nu$  mode.
- <sup>32</sup> ABULENCIA, A 06E measures the  $t\bar{t}$  production cross section in the all hadronic decay mode by selecting events with 6 to 8 jets and at least one  $b$ -jet.  $S/B = 1/5$  has been achieved. Based on  $311 \text{ pb}^{-1}$ . Assuming  $m_t = 178 \text{ GeV}$ .

- <sup>33</sup> Based on  $\sim 318 \text{ pb}^{-1}$ . Assuming  $m_t = 178 \text{ GeV}$ . Result is for at least one  $b$ -tag. For at least two  $b$ -tagged jets, the cross section is  $11.1^{+2.3+2.5}_{-1.9-1.9} \text{ pb}$ .
- <sup>34</sup> ABAZOV 05Q measures the top-quark pair production cross section with  $\sim 230 \text{ pb}^{-1}$  of data, based on the analysis of  $W$  plus  $n$ -jet events where  $W$  decays into  $e$  or  $\mu$  plus neutrino, and at least one of the jets is  $b$ -jet like. The first error is statistical and systematic, and the second accounts for the luminosity uncertainty. The result assumes  $m_t = 175 \text{ GeV}$ ; the mean value changes by  $(175 - m_t(\text{GeV})) \times 0.06 \text{ pb}$  in the mass range 160 to 190 GeV.
- <sup>35</sup> ABAZOV 05R measures the top-quark pair production cross section with  $224\text{--}243 \text{ pb}^{-1}$  of data, based on the analysis of events with two charged leptons in the final state. The result assumes  $m_t = 175 \text{ GeV}$ ; the mean value changes by  $(175 - m_t(\text{GeV})) \times 0.08 \text{ pb}$  in the mass range 160 to 190 GeV.
- <sup>36</sup> Based on  $230 \text{ pb}^{-1}$ . Assuming  $m_t = 175 \text{ GeV}$ .
- <sup>37</sup> Based on  $194 \text{ pb}^{-1}$ . Assuming  $m_t = 175 \text{ GeV}$ .
- <sup>38</sup> Based on  $194 \pm 11 \text{ pb}^{-1}$ . Assuming  $m_t = 175 \text{ GeV}$ .
- <sup>39</sup> Based on  $162 \pm 10 \text{ pb}^{-1}$ . Assuming  $m_t = 175 \text{ GeV}$ .
- <sup>40</sup> ACOSTA 05V measures the top-quark pair production cross section with  $\sim 162 \text{ pb}^{-1}$  of data, based on the analysis of  $W$  plus  $n$ -jet events where  $W$  decays into  $e$  or  $\mu$  plus neutrino, and at least one of the jets is  $b$ -jet like. Assumes  $m_t = 175 \text{ GeV}$ .
- <sup>41</sup> ACOSTA 04I measures the top-quark pair production cross section with  $197 \pm 12 \text{ pb}^{-1}$  of data, based on the analysis of events with two charged leptons in the final state. Assumes  $m_t = 175 \text{ GeV}$ .

## Ratio of the Production Cross Sections of $t\bar{t}\gamma$ to $t\bar{t}$ at $\sqrt{s} = 1.96 \text{ TeV}$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.024 \pm 0.009$	<sup>1</sup> AALTONEN 11Z	CDF	$E_T(\gamma) > 10 \text{ GeV}$ , $ \eta(\gamma)  < 1.0$
<sup>1</sup> Based on $6.0 \text{ fb}^{-1}$ of data. The error is statistical and systematic combined. Events with lepton + $E_T + \geq 3$ jets ( $\geq 1b$ ) with and without central, high $E_T$ photon are measured. The result is consistent with the SM prediction of $0.024 \pm 0.005$ . The absolute production cross section is measured to be $0.18 \pm 0.08 \text{ fb}$ . The statistical significance is 3.0 standard deviations.			

## $t\bar{t}$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 7 \text{ TeV}$

Unless otherwise noted the first quoted error is from statistics, the second from systematic uncertainties, and the third from luminosity. If only two errors are quoted the luminosity is included in the systematic uncertainties.

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$181.2 \pm 2.8^{+10.8}_{-10.6}$	<sup>1</sup> AAD	15B0 ATLAS	$e + \mu + E_T + \geq 0j$
$178 \pm 3 \pm 16 \pm 3$	<sup>2</sup> AAD <sup>3</sup> AAIJ	15CC ATLAS 15R LHCB	$\ell$ +jets, $\ell\bar{\ell}$ +jets, $\ell\tau_h$ +jets $\mu + \geq 1j$ ( $b$ -tag) forward region
$182.9 \pm 3.1 \pm 6.4$	<sup>4</sup> AAD	14AY ATLAS	$e + \mu + 1$ or 2b jets
$194 \pm 18 \pm 46$	<sup>5</sup> AAD	13X ATLAS	$\tau_h + E_T + \geq 5j$ ( $\geq 2b$ )
$139 \pm 10 \pm 26$	<sup>6</sup> CHATRCHYAN13AY	CMS	$\geq 6$ jets with 2 $b$ -tags
$158.1 \pm 2.1 \pm 10.8$	<sup>7</sup> CHATRCHYAN13BB	CMS	$\ell + E_T + \text{jets} (\geq 1 \text{ } b\text{-tag})$
$152 \pm 12 \pm 32$	<sup>8</sup> CHATRCHYAN13BE	CMS	$\tau_h + E_T + \geq 4 \text{ jets} (\geq 1 \text{ } b)$
$177 \pm 20 \pm 14 \pm 7$	<sup>9</sup> AAD	12B ATLAS	Repl. by AAD 12BF
$176 \pm 5 \pm 14 \pm 8$	<sup>10</sup> AAD	12BF ATLAS	$\ell\bar{\ell} + E_T + \geq 2j$
$187 \pm 11 \pm 18 \pm 6$	<sup>11</sup> AAD	12B0 ATLAS	$\ell + E_T + \geq 3j$ with $b$ -tag
$186 \pm 13 \pm 20 \pm 7$	<sup>12</sup> AAD	12CG ATLAS	$\ell + \tau_h + E_T + \geq 2j$ ( $\geq 1b$ )
$143 \pm 14 \pm 22 \pm 3$	<sup>13</sup> CHATRCHYAN12AC	CMS	$\ell + \tau_h + E_T + \geq 2j$ ( $\geq 1b$ )
$161.9 \pm 2.5^{+5.1}_{-5.0} \pm 3.6$	<sup>14</sup> CHATRCHYAN12AX	CMS	$\ell\bar{\ell} + E_T + \geq 2b$
$145 \pm 31 \pm 42 \pm 27$	<sup>15</sup> AAD	11A ATLAS	$\ell + E_T + \geq 4j$ , $\ell\bar{\ell} + E_T + \geq 2j$
$173 \pm 39 \pm 32 \pm 7$	<sup>16</sup> CHATRCHYAN11AA	CMS	$\ell + E_T + \geq 3 \text{ jets}$
$168 \pm 18 \pm 14 \pm 7$	<sup>17</sup> CHATRCHYAN11F	CMS	$\ell\bar{\ell} + E_T + \text{jets}$
$154 \pm 17 \pm 6$	<sup>18</sup> CHATRCHYAN11Z	CMS	Combination
$194 \pm 72 \pm 24 \pm 21$	<sup>19</sup> KHACHATRYAN11A	CMS	$\ell\bar{\ell} + E_T + \geq 2 \text{ jets}$

- <sup>1</sup> Based on  $4.6 \text{ fb}^{-1}$  of data. Uses a template fit to distributions of  $E_T$  and jet multiplicities to measure simultaneously  $t\bar{t}$ ,  $WW$ , and  $Z/\gamma^* \rightarrow \tau\tau$  cross sections, assuming  $m_t = 172.5 \text{ GeV}$ .
- <sup>2</sup> AAD 15CC based on  $4.6 \text{ fb}^{-1}$  of data. The event selection criteria are optimized for the  $\ell\tau_h + \text{jets}$  channel. Using only this channel  $183 \pm 9 \pm 23 \pm 3 \text{ pb}$  is derived for the cross section.
- <sup>3</sup> AAIJ 15R, based on  $1.0 \text{ fb}^{-1}$  of data, reports  $0.239 \pm 0.053 \pm 0.033 \pm 0.024 \text{ pb}$  cross section for the forward fiducial region  $p_T(\mu) > 25 \text{ GeV}$ ,  $2.0 < \eta(\mu) < 4.5$ ,  $50 \text{ GeV} < p_T(b) < 100 \text{ GeV}$ ,  $2.2 < \eta(b) < 4.2$ ,  $\Delta R(\mu, b) > 0.5$ , and  $p_T(\mu+b) > 20 \text{ GeV}$ . The three errors are from statistics, systematics, and theory. The result agrees with the SM NLO prediction.
- <sup>4</sup> AAD 14AY reports  $182.9 \pm 3.1 \pm 4.2 \pm 3.6 \pm 3.3 \text{ pb}$  value based on  $4.6 \text{ fb}^{-1}$  of data. The four errors are from statistics, systematic, luminosity, and the 0.66% beam energy uncertainty. We have combined the systematic uncertainties in quadrature. The result is for  $m_t = 172.5 \text{ GeV}$ ; for other  $m_t$ ,  $\sigma(m_t) = \sigma(172.5 \text{ GeV}) \times [1 - 0.0028 \times (m_t - 172.5 \text{ GeV})]$ . The result is consistent with the SM prediction at NNLO.
- <sup>5</sup> Based on  $1.67 \text{ fb}^{-1}$  of data. The result uses the acceptance for  $m_t = 172.5 \text{ GeV}$ .
- <sup>6</sup> Based on  $3.54 \text{ fb}^{-1}$  of data.
- <sup>7</sup> Based on  $2.3 \text{ fb}^{-1}$  of data.
- <sup>8</sup> Based on  $3.9 \text{ fb}^{-1}$  of data.
- <sup>9</sup> Based on  $35 \text{ pb}^{-1}$  of data for an assumed top quark mass of  $m_t = 172.5 \text{ GeV}$ .
- <sup>10</sup> Based on  $0.70 \text{ fb}^{-1}$  of data. The 3 errors are from statistics, systematics, and luminosity. The result uses the acceptance for  $m_t = 172.5 \text{ GeV}$ .
- <sup>11</sup> Based on  $35 \text{ pb}^{-1}$  of data. The 3 errors are from statistics, systematics, and luminosity. The result uses the acceptance for  $m_t = 172.5 \text{ GeV}$  and  $173 \pm 17^{+18}_{-16} \pm 6 \text{ pb}$  is found without the  $b$ -tag.

<sup>12</sup>Based on 2.05 fb<sup>-1</sup> of data. The hadronic  $\tau$  candidates are selected using a BDT technique. The 3 errors are from statistics, systematics, and luminosity. The result uses the acceptance for  $m_t = 172.5$  GeV.

<sup>13</sup>Based on 2.0 fb<sup>-1</sup> and 2.2 fb<sup>-1</sup> of data for  $\ell = e$  and  $\ell = \mu$ , respectively. The 3 errors are from statistics, systematics, and luminosity. The result uses the acceptance for  $m_t = 172.5$  GeV.

<sup>14</sup>Based on 2.3 fb<sup>-1</sup> of data. The 3 errors are from statistics, systematics, and luminosity. The result uses the profile likelihood-ratio (PLB) method and an assumed  $m_t$  of 172.5 GeV.

<sup>15</sup>Based on 2.9 pb<sup>-1</sup> of data. The result for single lepton channels is  $142 \pm 34^{+50}_{-31}$  pb, while for the dilepton channels is  $151^{+78+37}_{-62-24}$  pb.

<sup>16</sup>Result is based on 36 pb<sup>-1</sup> of data. The first uncertainty corresponds to the statistical and systematic uncertainties, and the second corresponds to the luminosity.

<sup>17</sup>Based on 36 pb<sup>-1</sup> of data. The ratio of  $t\bar{t}$  and  $Z/\gamma^*$  cross sections is measured as  $\sigma(pp \rightarrow t\bar{t})/\sigma(pp \rightarrow Z/\gamma^* \rightarrow e^+e^-/\mu^+\mu^-) = 0.175 \pm 0.018(\text{stat}) \pm 0.015(\text{syst})$  for  $60 < m_{\ell\ell} < 120$  GeV, for which they use an NNLO prediction for the denominator cross section of 972  $\pm$  42 pb.

<sup>18</sup>Result is based on 36 pb<sup>-1</sup> of data. The first error is from statistical and systematic uncertainties, and the second from luminosity. This is a combination of a measurement in the dilepton channel (CHATRCHYAN 11f) and the measurement in the  $\ell$  + jets channel (CHATRCHYAN 11z) which yields  $150 \pm 9 \pm 17 \pm 6$  pb.

<sup>19</sup>Result is based on  $3.1 \pm 0.3$  pb<sup>-1</sup> of data.

### $t\bar{t}$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 8$ TeV

Unless otherwise noted the first quoted error is from statistics, the second from systematic uncertainties, and the third from luminosity. If only two errors are quoted the luminosity is included in the systematic uncertainties.

VALUE (pb)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
260 $\pm 1$ $^{+24}_{-25}$	<sup>1</sup> AAD	15BP ATLS	$\ell + \cancel{E}_T + \geq 3j$ ( $\geq 1b$ )
	<sup>2</sup> AAIJ	15R LHCb	$\mu + \geq 1j(b\text{-tag})$ forward region
242.4 $\pm 1.7 \pm 10.2$	<sup>3</sup> AAD	14AY ATLS	$e + \mu + 1$ or $2b$ jets
239 $\pm 2$ $\pm 11 \pm 6$	<sup>4</sup> CHATRCHYAN14f	CMS	$\ell\ell + \cancel{E}_T + \geq 2j$ ( $\geq 1 b\text{-tag}$ )
257 $\pm 3$ $\pm 24 \pm 7$	<sup>5</sup> KHACHATRY...14s	CMS	$\ell + \tau_h + \cancel{E}_T + \geq 2j$ ( $\geq 1b$ )

<sup>1</sup>AAD 15BP based on 20.3 fb<sup>-1</sup> of data. The result is for  $m_t = 172.5$  GeV and in agreement with the SM prediction  $253^{+13}_{-15}$  pb at NNLO+NNLL.

<sup>2</sup>AAIJ 15R, based on 2.0 fb<sup>-1</sup> of data, reports  $0.289 \pm 0.043 \pm 0.040 \pm 0.029$  pb cross section for the forward fiducial region  $p_T(\mu) > 25$  GeV,  $2.0 < \eta(\mu) < 4.5$ ,  $50 \text{ GeV} < p_T(b) < 100$  GeV,  $2.2 < \eta(b) < 4.2$ ,  $\Delta R(\mu, b) > 0.5$ , and  $p_T(\mu + b) > 20$  GeV. The three errors are from statistics, systematics, and theory. The result agrees with the SM NLO prediction.

<sup>3</sup>AAD 14AY reports  $242.4 \pm 1.7 \pm 5.5 \pm 7.5 \pm 4.2$  pb value based on 20.3 fb<sup>-1</sup> of data. The four errors are from statistics, systematic, luminosity, and the 0.66% beam energy uncertainty. We have combined the systematic uncertainties in quadrature. The result is for  $m_t = 172.5$  GeV; for other  $m_t$ ,  $\sigma(m_t) = \sigma(172.5 \text{ GeV}) \times [1 - 0.0028 \times (m_t - 172.5 \text{ GeV})]$ . Also measured is the ratio  $\sigma(t\bar{t}; 8 \text{ TeV})/\sigma(t\bar{t}; 7 \text{ TeV}) = 1.326 \pm 0.024 \pm 0.015 \pm 0.049 \pm 0.001$ . The results are consistent with the SM predictions at NNLO.

<sup>4</sup>Based on 5.3 fb<sup>-1</sup> of data. The result is for  $m_t = 172.5$  GeV, and a parametrization is given in eq.(6.1) for the mean value at other  $m_t$  values. The result is in agreement with the SM prediction  $252.9^{+6.4}_{-8.6}$  pb at NNLO.

<sup>5</sup>Based on 19.6 fb<sup>-1</sup> of data. The measurement is in the channel  $t\bar{t} \rightarrow (b\ell\nu)(b\tau\nu)$ , where  $\tau$  decays into hadrons ( $\tau_h$ ). The result is for  $m_t = 172.5$  GeV. For  $m_t = 173.3$  GeV, the cross section is lower by 3.1 pb.

### $t t$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 7$ TeV

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
<1.7	95	<sup>1</sup> AAD	12BE ATLS	$\ell + e^+ + \cancel{E}_T + \geq 2j + \text{HT}$

<sup>1</sup>Based on 1.04 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. The upper bounds are the same for LL, LR and RR chiral components of the two top quarks.

### $t\bar{t} t\bar{t}$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 8$ TeV

VALUE (fb)	CL%	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
<23	95	<sup>1</sup> AAD	15AR ATLS	$\ell + \cancel{E}_T + \geq 5j$ ( $\geq 2 b$ )
<70	95	<sup>2</sup> AAD	15BY ATLS	$\geq 2\ell + \cancel{E}_T + \geq 2j$ ( $\geq 1 b$ )
<32	95	<sup>3</sup> KHACHATRY...14r	CMS	$\ell + \cancel{E}_T + \geq 6j$ ( $\geq 2 b$ )

<sup>1</sup>AAD 15AR based on 20.3 fb<sup>-1</sup> of data. A fit to  $H_T$  distributions in multi-channels classified by the number of jets and of  $b$ -tagged jets is performed.

<sup>2</sup>AAD 15BY based on 20.3 fb<sup>-1</sup> of data. A same-sign lepton pair is required. An excess over the SM prediction reaches  $2.5\sigma$  for hypotheses involving heavy resonances decaying into  $t\bar{t}t\bar{t}$ .

<sup>3</sup>Based on 19.6 fb<sup>-1</sup> of data, using a multivariate analysis to separate signal from backgrounds. About  $\sigma(t\bar{t}t\bar{t}) = 1$  fb is expected in the SM.

### $t\bar{t} W$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 8$ TeV

VALUE (fb)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
170 $^{+90}_{-80} \pm 70$	<sup>1</sup> KHACHATRY...14N	CMS	$t\bar{t} W \rightarrow$ same sign dilepton + $\cancel{E}_T$ + jets

<sup>1</sup>Based on 19.5 fb<sup>-1</sup> of data. The result is consistent with the SM prediction of  $\sigma(t\bar{t} W) = 206^{+21}_{-23}$  fb.

### $t\bar{t} Z$ Production Cross Section in $pp$ Collisions at $\sqrt{s} = 8$ TeV

VALUE (fb)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
200 $^{+80+40}_{-70-30}$	<sup>1</sup> KHACHATRY...14N	CMS	$t\bar{t} Z \rightarrow 3, 4 \ell + \cancel{E}_T$ + jets

<sup>1</sup>Based on 19.5 fb<sup>-1</sup> of data. The result is consistent with the SM prediction of  $\sigma(t\bar{t} Z) = 197^{+22}_{-25}$  fb.

### $f(Q_0)$ : $t\bar{t}$ Fraction of Events with a Veto on Additional Central Jet Activity in $pp$ Collisions at $\sqrt{s} = 7$ TeV

$Q_0$  denotes the threshold of the additional jet  $p_T$ .

VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
80.0 $\pm 1.1 \pm 1.6$	<sup>1</sup> CHATRCHYAN14AE	CMS	$Q_0 = 75 \text{ GeV}$ ( $ \eta  < 2.4$ )
92.0 $\pm 0.7 \pm 0.8$	<sup>1</sup> CHATRCHYAN14AE	CMS	$Q_0 = 150 \text{ GeV}$ ( $ \eta  < 2.4$ )
98.0 $\pm 0.3 \pm 0.3$	<sup>1</sup> CHATRCHYAN14AE	CMS	$Q_0 = 300 \text{ GeV}$ ( $ \eta  < 2.4$ )
56.4 $\pm 1.3^{+2.6}_{-2.8}$	<sup>2</sup> AAD	12BL ATLS	$Q_0 = 25 \text{ GeV}$ ( $ \eta  < 2.1$ )
84.7 $\pm 0.9 \pm 1.0$	<sup>2</sup> AAD	12BL ATLS	$Q_0 = 75 \text{ GeV}$ ( $ \eta  < 2.1$ )
95.2 $^{+0.5}_{-0.6} \pm 0.4$	<sup>2</sup> AAD	12BL ATLS	$Q_0 = 150 \text{ GeV}$ ( $ \eta  < 2.1$ )

<sup>1</sup>CHATRCHYAN 15 based on 5.0 fb<sup>-1</sup> of data. The  $t\bar{t}$  events are selected in the dilepton and lepton + jets decay channels. For other values of  $Q_0$  see Table 5.

<sup>2</sup>Based on 2.05 fb<sup>-1</sup> of data. The  $t\bar{t}$  events are selected in the dilepton decay channel with two identified  $b$ -jets.

### Fraction of $t\bar{t}$ + multi-jet Events in $pp$ Collisions at $\sqrt{s} = 7$ TeV

VALUE	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
	<sup>1</sup> AAD	15D ATLS	$\ell + \cancel{E}_T + n_j$ ( $n=3$ to 8)
0.332 $\pm 0.090$	<sup>2</sup> CHATRCHYAN14AE	CMS	$t\bar{t}(\ell\ell) + 0 \text{ jet}$ ( $E_T > 30 \text{ GeV}$ )
0.436 $\pm 0.098$	<sup>2</sup> CHATRCHYAN14AE	CMS	$t\bar{t}(\ell\ell) + 1 \text{ jet}$ ( $E_T > 30 \text{ GeV}$ )
0.232 $\pm 0.125$	<sup>2</sup> CHATRCHYAN14AE	CMS	$t\bar{t}(\ell\ell) + \geq 2 \text{ jet}$ ( $E_T > 30 \text{ GeV}$ )

<sup>1</sup>Based on 4.6 fb<sup>-1</sup> of data. Fiducial  $t\bar{t}$  production cross section is presented as a function of the jet multiplicity for up to eight jets with the jet  $p_T$  threshold of 25, 40, 60, and 80 GeV, and as a function of jet  $p_T$  up to the 5th jet. MC models can be discriminated by using data for high jet multiplicity and by  $p_T$  distributions of the leading and 5th jet.

<sup>2</sup>Based on 5.0 fb<sup>-1</sup> of data. Events with two oppositely charged leptons, large  $\cancel{E}_T$  and jets with at least 1  $b$ -tag are used to measure the fraction of  $t\bar{t}$  plus additional jets. The gap fraction ( $n=0$  jet rate) as a function of the jet  $p_T$  and that of  $H_T$ , the scalar sum of the  $p_T$ 's of additional jets, is shown in Fig. 8.

### $t\bar{t}$ Charge Asymmetry ( $A_C$ ) in $pp$ Collisions at $\sqrt{s} = 7$ TeV

$A_C = (N(\Delta|y| > 0) - N(\Delta|y| < 0)) / (N(\Delta|y| > 0) + N(\Delta|y| < 0))$  where  $\Delta|y| = |y_t| - |y_{\bar{t}}|$  is the difference between the absolute values of the top and antitop rapidities and  $N$  is the number of events with  $\Delta|y|$  positive or negative.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
2.1 $\pm 2.5 \pm 1.7$	<sup>1</sup> AAD	15AJ ATLS	$\ell\ell + \cancel{E}_T + \geq 2j$
0.6 $\pm 1.0$	<sup>2</sup> AAD	14i ATLS	$\ell + \cancel{E}_T + \geq 4j$ ( $\geq 1b$ )
-1.0 $\pm 1.7 \pm 0.8$	<sup>3</sup> CHATRCHYAN14d	CMS	$\ell\ell + \cancel{E}_T + \geq 2j$ ( $\geq 1b$ )
-1.9 $\pm 2.8 \pm 2.4$	<sup>4</sup> AAD	12BK ATLS	$\ell + \cancel{E}_T + \geq 4j$ ( $\geq 1b$ )
0.4 $\pm 1.0 \pm 1.1$	<sup>5</sup> CHATRCHYAN12BB	CMS	$\ell + \cancel{E}_T + \geq 4j$ ( $\geq 1b$ )
-1.3 $\pm 2.8 \pm 2.9$ $^{+2.9}_{-3.1}$	<sup>6</sup> CHATRCHYAN12Bs	CMS	$\ell + \cancel{E}_T + \geq 4j$ ( $\geq 1b$ )

<sup>1</sup>AAD 15AJ based on 4.6 fb<sup>-1</sup> of data. After kinematic reconstruction the top quark momenta are corrected for detector resolution and acceptance effects by unfolding, using parton level information of the MC generators. The lepton charge asymmetry is measured as  $A_C^\ell = 0.024 \pm 0.015 \pm 0.009$ . All the measurements are consistent with the SM predictions.

<sup>2</sup>Based on 4.7 fb<sup>-1</sup> of data. The result is consistent with the SM prediction of  $A_C = 0.0123 \pm 0.0005$ . The asymmetry is  $0.011 \pm 0.018$  if restricted to those events where  $\beta_Z(t\bar{t}) > 0.6$ , which is also consistent with the SM prediction of  $0.020^{+0.006}_{-0.007}$ .

<sup>3</sup>Based on 5.0 fb<sup>-1</sup> of data. The lepton charge asymmetry is measured as  $A_C^\ell = 0.009 \pm 0.0010 \pm 0.006$ .  $A_C^\ell$  dependences on  $m_{t\bar{t}}$ ,  $|y(t\bar{t})|$ , and  $p_T(t\bar{t})$  are given in Fig. 5. All measurements are consistent with the SM predictions.

<sup>4</sup>Based on 1.04 fb<sup>-1</sup> of data. The result is consistent with  $A_C = 0.006 \pm 0.002$  (MC at NLO). No significant dependence of  $A_C$  on  $m_{t\bar{t}}$  is observed.

<sup>5</sup>Based on 5.0 fb<sup>-1</sup> of data at 7 TeV.

<sup>6</sup>Based on 1.09 fb<sup>-1</sup> of data. The result is consistent with the SM predictions.

### $t$ -quark Polarization in $t\bar{t}$ Events in $p\bar{p}$ Collisions at $\sqrt{s} = 1.96$ TeV

VALUE	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
0.113 $\pm 0.091 \pm 0.019$	<sup>1</sup> ABAZOV	15K D0	$A_{FB}^\ell$ in $\ell\ell + \cancel{E}_T + \geq 2j$ ( $\geq 1b$ )

<sup>1</sup>ABAZOV 15K based on 9.7 fb<sup>-1</sup> of data. The value is top quark polarization times spin analyzing power in the beam basis. The result is consistent with the SM prediction of  $-0.0019 \pm 0.0005$ .

# Quark Particle Listings

*t*

## *t*-quark Polarization in *t* $\bar{t}$ Events in *p**p* Collisions at $\sqrt{s} = 7$ TeV

The double differential distribution in polar angles,  $\theta_1$  ( $\theta_2$ ) of the decay particle of the top (anti-top) decay products, is parametrized as  $(1/\sigma)d\sigma/(d\cos\theta_1 d\cos\theta_2) = (1/4) (1 + A_+ \cos\theta_1 + A_- \cos\theta_2 - C \cos\theta_1 \cos\theta_2)$ . The charged lepton is used to tag *t* or  $\bar{t}$ . The coefficient  $A_+$  and  $A_-$  measure the average helicity of *t* and  $\bar{t}$ , respectively.  $A_{CPC}$  assumes *CP* conservation, whereas  $A_{CPV}$  corresponds to maximal *CP* violation.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.035 \pm 0.014 \pm 0.037$	<sup>1</sup> AAD	13BE ATLS	$A_{CPC} = A_t = A_{\bar{t}}$
$0.020 \pm 0.016^{+0.013}_{-0.017}$	<sup>1</sup> AAD	13BE ATLS	$A_{CPV} = A_t = -A_{\bar{t}}$

<sup>1</sup> Based on 4.7 fb<sup>-1</sup> of data using the final states containing one or two isolated electrons or muons and jets with at least one *b*-tag.

## *g**g* → *t* $\bar{t}$ Fraction in *p* $\bar{p}$ Collisions at $\sqrt{s} = 1.96$ TeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.33	68	<sup>1</sup> AALTONEN	09F CDF	<i>t</i> $\bar{t}$ correlations
$0.07 \pm 0.14 \pm 0.07$		<sup>2</sup> AALTONEN	08AG CDF	low $p_T$ number of tracks

<sup>1</sup> Based on 955 pb<sup>-1</sup>. AALTONEN 09F used differences in the *t* $\bar{t}$  production angular distribution and polarization correlation to discriminate between *g**g* → *t* $\bar{t}$  and  $q\bar{q}$  → *t* $\bar{t}$  subprocesses. The combination with the result of AALTONEN 08AG gives  $0.07^{+0.15}_{-0.07}$ .

<sup>2</sup> Result is based on 0.96 fb<sup>-1</sup> of data. The contribution of the subprocesses *g**g* → *t* $\bar{t}$  and  $q\bar{q}$  → *t* $\bar{t}$  is distinguished by using the difference between quark and gluon initiated jets in the number of small  $p_T$  (0.3 GeV <  $p_T$  < 3 GeV) charged particles in the central region ( $|\eta| < 1.1$ ).

## $A_{FB}^{\ell}$ of *t* $\bar{t}$ in *p* $\bar{p}$ Collisions at $\sqrt{s} = 1.96$ TeV

VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$17.5 \pm 5.6 \pm 3.1$	<sup>1</sup> ABAZOV	15K D0	$A_{FB}^{\ell}$ in $\ell\ell + \cancel{E}_T + \geq 2j$ ( $\geq 1b$ )
$7.2 \pm 6.0$	<sup>2</sup> AALTONEN	14F CDF	$A_{FB}^{\ell}$ in dilepton channel ( $\ell\ell + \cancel{E}_T + \geq 2j$ )
$7.6 \pm 8.2$	<sup>2</sup> AALTONEN	14F CDF	$A_{FB}^{\ell\ell}$ in dilepton channel ( $\ell\ell + \cancel{E}_T + \geq 2j$ )
$4.2 \pm 2.3^{+1.7}_{-2.0}$	<sup>3</sup> ABAZOV	14G D0	$A_{FB}^{\ell}$ ( $\ell + \cancel{E}_T + \geq 3j$ (0.1 $\geq 2b$ ))
$10.6 \pm 3.0$	<sup>4</sup> ABAZOV	14H D0	$A_{FB}^{\ell}$ ( $\ell + \cancel{E}_T + \geq 3j$ ( $\geq 1b$ ))
$20.1 \pm 6.7$	<sup>5</sup> AALTONEN	13AD CDF	$a_1/a_0$ in $\ell + \cancel{E}_T + \geq 4j$ ( $\geq 1b$ )
$-0.2 \pm 3.1$	<sup>5</sup> AALTONEN	13AD CDF	$a_3, a_5, a_7$ in $\ell + \cancel{E}_T + \geq 4j$ ( $\geq 1b$ )
$16.4 \pm 4.7$	<sup>6</sup> AALTONEN	13S CDF	$\ell + \cancel{E}_T + \geq 4$ jets ( $\geq 1b$ -tag)
$9.4^{+3.2}_{-2.9}$	<sup>7</sup> AALTONEN	13X CDF	$\ell + \cancel{E}_T + \geq 4$ jets ( $\geq 1$ <i>b</i> -tag)
$11.8 \pm 3.2$	<sup>8</sup> ABAZOV	13A D0	$\ell\ell$ & $\ell+$ jets comb.
$-11.6 \pm 15.3$	<sup>9</sup> AALTONEN	11F CDF	$m_{t\bar{t}} < 450$ GeV
$47.5 \pm 11.4$	<sup>9</sup> AALTONEN	11F CDF	$m_{t\bar{t}} > 450$ GeV
$19.6 \pm 6.5$	<sup>10</sup> ABAZOV	11AH D0	$\ell + \cancel{E}_T + \geq 4$ jets ( $\geq 1b$ -tag)
$17 \pm 8$	<sup>11</sup> AALTONEN	08AB CDF	<i>p</i> $\bar{p}$ frame
$24 \pm 14$	<sup>11</sup> AALTONEN	08AB CDF	<i>t</i> $\bar{t}$ frame
$12 \pm 8 \pm 1$	<sup>12</sup> ABAZOV	08L D0	$\ell + \cancel{E}_T + \geq 4$ jets

<sup>1</sup> ABAZOV 15K based on 9.7 fb<sup>-1</sup> of data. The result is consistent with the SM predictions. By combining with the previous D0 measurement in the  $\ell$  + jet channel ABAZOV 14H,  $A_{FB}^{\ell} = 0.118 \pm 0.025 \pm 0.013$  is obtained.

<sup>2</sup> Based on 9.1 fb<sup>-1</sup> of data. Both results are consistent with the SM predictions. By combining with the previous CDF measurement in the  $\ell$  + jet channel AALTONEN 13X,  $A_{FB}^{\ell} = 0.090^{+0.028}_{-0.026}$  is obtained. The combined result is about two sigma larger than the SM prediction of  $A_{FB}^{\ell} = 0.038 \pm 0.003$ .

<sup>3</sup> Based on 9.7 fb<sup>-1</sup> of *p* $\bar{p}$  data at  $\sqrt{s} = 1.96$  TeV. The asymmetry is corrected for the production level for events with  $|y_{\ell}| < 1.5$ . Asymmetry as functions of  $E_T(\ell)$  and  $|y_{\ell}|$  are given in Figs. 7 and 8, respectively. Combination with the asymmetry measured in the dilepton channel [ABAZOV 13P] gives  $A_{FB}^{\ell} = 4.2 \pm 2.0 \pm 1.4$  %, in agreement with the SM prediction of 2.0%.

<sup>4</sup> Based on 9.7 fb<sup>-1</sup> of data of *p* $\bar{p}$  data at  $\sqrt{s} = 1.96$  TeV. The measured asymmetry is in agreement with the SM predictions of  $8.8 \pm 0.9$  % [BERNREUTHER 12], which includes the EW effects. The dependences of the asymmetry on  $|y(t) - y(\bar{t})|$  and  $m_{t\bar{t}}$  are shown in Figs. 9 and 10, respectively.

<sup>5</sup> Based on 9.4 fb<sup>-1</sup> of data. Reported  $A_{FB}^{\ell}$  values come from the determination of  $a_i$  coefficients of  $d\sigma/d(\cos\theta_{\ell}) = \sum_i a_i P_i(\cos\theta_{\ell})$  measurement. The result of  $a_1/a_0 = (40 \pm 12)\%$  seems higher than the NLO SM prediction of  $(15^{+7}_{-3})\%$ .

<sup>6</sup> Based on 9.4 fb<sup>-1</sup> of data. The quoted result is the asymmetry at the parton level.

<sup>7</sup> Based on 9.4 fb<sup>-1</sup> of data. The observed asymmetry is to be compared with the SM prediction of  $A_{FB}^{\ell} = 0.038 \pm 0.003$ .

<sup>8</sup> Based on 5.4 fb<sup>-1</sup> of data. ABAZOV 13A studied the dilepton channel of the *t* $\bar{t}$  events and measured the leptonic forward-backward asymmetry to be  $A_{FB}^{\ell} = 5.8 \pm 5.1 \pm 1.3\%$ , which is consistent with the SM (QCD+EW) prediction of  $4.7 \pm 0.1\%$ . The result is obtained after combining the measurement  $(15.2 \pm 4.0\%)$  in the  $\ell$  + jets channel ABAZOV 11AH. The top quark helicity is measured by using the neutrino weighting method to be consistent with zero in both dilepton and  $\ell$  + jets channels.

<sup>9</sup> Based on 5.3 fb<sup>-1</sup> of data. The error is statistical and systematic combined. Events with lepton +  $E_T + \geq 4$  jets ( $\geq 1b$ ) are used. AALTONEN 11F also measures the asymmetry as a function of the rapidity difference  $|y_{\ell} - y_{\bar{\ell}}|$ . The NLO QCD predictions [MCFM] are  $(4.0 \pm 0.6)\%$  and  $(8.8 \pm 1.3)\%$  for  $m_{t\bar{t}} < 450$  and  $> 450$  GeV, respectively.

<sup>10</sup> Based on 5.4 fb<sup>-1</sup> of data. The error is statistical and systematic combined. The quoted asymmetry is obtained after unfolding to be compared with the MC@NLO prediction of  $(5.0 \pm 0.1)\%$ . No significant difference between the  $m_{t\bar{t}} < 450$  and  $> 450$  GeV data samples is found. A corrected asymmetry based on the lepton from a top quark decay of  $(15.2 \pm 4.0)\%$  is measured to be compared to the MC@NLO prediction of  $(2.1 \pm 0.1)\%$ .

<sup>11</sup> Result is based on 1.9 fb<sup>-1</sup> of data. The *FB* asymmetry in the *t* $\bar{t}$  events has been measured in the  $\ell$  + jets mode, where the lepton charge is used as the flavor tag. The asymmetry in the *p* $\bar{p}$  frame is defined in terms of  $\cos(\theta)$  of hadronically decaying *t*-quark momentum, whereas that in the *t* $\bar{t}$  frame is defined in terms of the *t* and  $\bar{t}$  rapidity difference. The results are consistent ( $\leq 2\sigma$ ) with the SM predictions.

<sup>12</sup> Result is based on 0.9 fb<sup>-1</sup> of data. The asymmetry in the number of *t* $\bar{t}$  events with  $y_t > y_{\bar{t}}$  and those with  $y_t < y_{\bar{t}}$  has been measured in the lepton + jets final state. The observed value is consistent with the SM prediction of 0.8% by MC@NLO, and an upper bound on the  $Z' \rightarrow t\bar{t}$  contribution for the SM Z-like couplings is given in Fig. 2 for  $350 \text{ GeV} < m_{Z'} < 1 \text{ TeV}$ .

## *t*-Quark Electric Charge

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.64 \pm 0.02 \pm 0.08</math></b>	<sup>1</sup> AAD	13AY ATLS	$\ell + \cancel{E}_T + \geq 4$ jets ( $\geq 1$ <i>b</i> )
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	<sup>2</sup> ABAZOV	14D D0	$\ell + \cancel{E}_T + \geq 4$ jets ( $\geq 2$ <i>b</i> )
	<sup>3</sup> AALTONEN	13J CDF	<i>p</i> $\bar{p}$ at 1.96 TeV
	<sup>4</sup> AALTONEN	10S CDF	Repl. by AALTONEN 13J
	<sup>5</sup> ABAZOV	07c D0	fraction of $ q  = 4e/3$ pair

<sup>1</sup> AAD 13AY result is based on 2.05 fb<sup>-1</sup> of *p**p* data at  $\sqrt{s} = 7$  TeV, the result is obtained by reconstructing *t* $\bar{t}$  events in the lepton + jets final state, where *b*-jet charges are tagged by the jet-charge algorithm. This measurement excludes the charge  $-4/3$  assignment to the top quark at more than 8 standard deviations.

<sup>2</sup> ABAZOV 14D result is based on 5.3 fb<sup>-1</sup> of *p* $\bar{p}$  data at  $\sqrt{s} = 1.96$  TeV. The electric charge of *b* + *W* system in *t* $\bar{t}$  candidate events is measured from the charges of the leptons from *W* decay and in *b* jets. Under the assumption that the *b* + *W* system consists of the sum of the top quark and the charge  $-4/3$  quark  $b'(-4/3)$  of the same mass, the top quark fraction is found to be  $f = 0.88 \pm 0.13$  (stat)  $\pm 0.11$  (syst), or the upper bound for the  $b'(-4/3)$  contamination of  $1 - f < 0.46$  (95% CL).

<sup>3</sup> AALTONEN 13J excludes the charge  $-4/3$  assignment to the top quark at 99% CL, using 5.6 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. Result is obtained by reconstructing *t* $\bar{t}$  events in the lepton + jets final state, where *b*-jet charges are tagged by the jet-charge algorithm.

<sup>4</sup> AALTONEN 10S excludes the charge  $-4/3$  assignment for the top quark [CHANG 99] at 95% CL, using 2.7 fb<sup>-1</sup> of data in *p* $\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. Result is obtained by reconstructing *t* $\bar{t}$  events in the lepton + jets final state, where *b*-jet charges are tagged by the SLT (soft lepton tag) algorithm.

<sup>5</sup> ABAZOV 07c reports an upper limit  $\rho < 0.80$  (90% CL) on the fraction  $\rho$  of exotic quark pairs  $Q\bar{Q}$  with electric charge  $|q| = 4e/3$  in *t* $\bar{t}$  candidate events with high  $p_T$  lepton, missing  $E_T$  and  $\geq 4$  jets. The result is obtained by measuring the fraction of events in which the quark pair decays into  $W^- + b$  and  $W^+ + \bar{b}$ , where *b* and  $\bar{b}$  jets are discriminated by using the charge and momenta of tracks within the jet cones. The maximum CL at which the model of CHANG 99 can be excluded is 92%. Based on 370 pb<sup>-1</sup> of data at  $\sqrt{s} = 1.96$  TeV.

## *t*-Quark REFERENCES

AAD	16B	JHEP 1601 064	G. Aad et al.	(ATLAS Collab.)
AAD	16D	EPJ C76 12	G. Aad et al.	(ATLAS Collab.)
AALTONEN	16	PR D93 032011	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	16	PL B752 18	V.M. Abazov et al.	(D0 Collab.)
AAD	15	PL B740 222	G. Aad et al.	(ATLAS Collab.)
AAD	15A	PL B740 118	G. Aad et al.	(ATLAS Collab.)
AAD	15AJ	JHEP 1505 061	G. Aad et al.	(ATLAS Collab.)
AAD	15AR	JHEP 1508 105	G. Aad et al.	(ATLAS Collab.)
AAD	15AW	EPJ C75 158	G. Aad et al.	(ATLAS Collab.)
AAD	15BF	EPJ C75 330	G. Aad et al.	(ATLAS Collab.)
AAD	15BQ	PR D91 052005	G. Aad et al.	(ATLAS Collab.)
AAD	15BP	PR D91 112013	G. Aad et al.	(ATLAS Collab.)
AAD	15BW	JHEP 1510 121	G. Aad et al.	(ATLAS Collab.)
AAD	15BY	JHEP 1510 150	G. Aad et al.	(ATLAS Collab.)
AAD	15CC	PR D92 072005	G. Aad et al.	(ATLAS Collab.)
AAD	15CO	JHEP 1512 061	G. Aad et al.	(ATLAS Collab.)
AAD	15D	JHEP 1501 020	G. Aad et al.	(ATLAS Collab.)
AAD	15J	PRL 114 142001	G. Aad et al.	(ATLAS Collab.)
AAD	15R	PRL 115 112001	R. Aaij et al.	(LHCb Collab.)
AALTONEN	15D	PR D92 032003	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	15H	PRL 115 152003	T. Aaltonen et al.	(CDF, DO Collab.)
ABAZOV	15G	PR D91 112003	V.M. Abazov et al.	(DO Collab.)
ABAZOV	15K	PR D92 052007	V.M. Abazov et al.	(DO Collab.)
CHATRCHYAN	15	EPJ C75 216 (err.)	S. Chatrchyan et al.	(CMS Collab.)
AAD	14	PL B728 363	G. Aad et al.	(ATLAS Collab.)
AAD	14AA	JHEP 1406 008	G. Aad et al.	(ATLAS Collab.)
AAD	14AY	EPJ C74 3109	G. Aad et al.	(ATLAS Collab.)
AAD	14BB	PR D90 112016	G. Aad et al.	(ATLAS Collab.)
AAD	14BI	PR D90 112006	G. Aad et al.	(ATLAS Collab.)
AAD	14I	JHEP 1402 107	G. Aad et al.	(ATLAS Collab.)
AALTONEN	14F	PR D89 091101	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	14F	PRL 113 042001	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	14G	PRL 112 221801	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	14H	PR D89 072001	T. Aaltonen et al.	(CDF, DO Collab.)
AALTONEN	14K	PRL 112 231805	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	14L	PRL 112 231804	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	14M	PRL 112 231803	T. Aaltonen et al.	(CDF, DO Collab.)
AALTONEN	14N	PR D90 091101	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	14O	PRL 113 261804	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	14C	PRL 113 032002	V.M. Abazov et al.	(DO Collab.)
Also		PR D91 112003	V.M. Abazov et al.	(DO Collab.)
ABAZOV	14D	PR D90 051101	V.M. Abazov et al.	(DO Collab.)
ABAZOV	14G	PR D90 072001	V.M. Abazov et al.	(DO Collab.)
ABAZOV	14H	PR D90 072011	V.M. Abazov et al.	(DO Collab.)
ABAZOV	14K	PR D90 092006	V.M. Abazov et al.	(DO Collab.)
CHATRCHYAN	14	PL B728 496	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14AC	PRL 112 231802	S. Chatrchyan et al.	(CMS Collab.)



# Quark Particle Listings

## $t$ , $b'$ (Fourth Generation) Quark

ABE	95F	PRL 74 2626	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	94E	PR D50 2966	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PRL 73 225	F. Abe <i>et al.</i>	(CDF Collab.)

### $b'$ ( $4^{th}$ Generation) Quark, Searches for

#### $b'(-1/3)$ -quark/hadron mass limits in $p\bar{p}$ and $pp$ collisions

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>620	95	1 AAD	15BY ATLS	$Wt$ , $Zb$ , $hb$ modes
>730	95	2 AAD	15BY ATLS	$B(b' \rightarrow Wt) = 1$
>810	95	3 AAD	15Z ATLS	
>755	95	4 AAD	14AZ ATLS	
>675	95	5 CHATRCHYAN 13i	CMS	$B(b' \rightarrow Wt) = 1$
>190	95	6 ABAZOV	08x D0	$c\tau = 200\text{mm}$
>190	95	7 ACOSTA	03 CDF	quasi-stable $b'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<350, 580–635, >700	95	8 AAD	15AR ATLS	$B(b' \rightarrow Hb) = 1$
>690	95	9 AAD	15CN ATLS	$B(b' \rightarrow Wq) = 1$ ( $q=u$ )
>480	95	10 AAD	12AT ATLS	$B(b' \rightarrow Wt) = 1$
>400	95	11 AAD	12AU ATLS	$B(b' \rightarrow Zb) = 1$
>350	95	12 AAD	12bC ATLS	$B(b' \rightarrow Wq) = 1$ ( $q=u,c$ )
>450	95	13 AAD	12bE ATLS	$B(b' \rightarrow Wt) = 1$
>685	95	14 CHATRCHYAN 12bH	CMS	$m_{b'} > m_{b'}$
>611	95	15 CHATRCHYAN 12x	CMS	$B(b' \rightarrow Wt) = 1$
>372	95	16 AALTONEN	11j CDF	$b' \rightarrow Wt$
>361	95	17 CHATRCHYAN 11L	CMS	Repl. by CHATRCHYAN 12x
>338	95	18 AALTONEN	10H CDF	$b' \rightarrow Wt$
>380–430	95	19 FLACCO	10 RVUE	$m_{b'} > m_{b'}$
>268	95	20,21 AALTONEN	07c CDF	$B(b' \rightarrow Zb) = 1$
>199	95	22 AFFOLDER	00 CDF	NC: $b' \rightarrow Zb$
>148	95	23 ABE	98N CDF	NC: $b' \rightarrow Zb + \text{vertex}$
>96	95	24 ABACHI	97D D0	NC: $b' \rightarrow b\gamma$
>128	95	25 ABACHI	95F D0	$\ell\ell + \text{jets}$ , $\ell + \text{jets}$
>75	95	26 MUKHOPAD..	93 RVUE	NC: $b' \rightarrow b\ell\ell$
>85	95	27 ABE	92 CDF	CC: $\ell\ell$
>72	95	28 ABE	90B CDF	CC: $e + \mu$
>54	95	29 AKESSON	90 UA2	CC: $e + \text{jets} + \cancel{E}_T$
>43	95	30 ALBAJAR	90B UA1	CC: $\mu + \text{jets}$
>34	95	31 ALBAJAR	88 UA1	CC: $e$ or $\mu + \text{jets}$

1 AAD 15BY based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Limit on pair-produced vector-like  $b'$  assuming the branching fractions to  $W$ ,  $Z$ , and  $h$  modes of the singlet model. Used events containing  $\geq 2\ell + \cancel{E}_T + \geq 2j$  ( $\geq 1b$ ) and including a same-sign lepton pair.

2 AAD 15BY based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Limit on pair-produced chiral  $b'$ -quark. Used events containing  $\geq 2\ell + \cancel{E}_T + \geq 2j$  ( $\geq 1b$ ) and including a same-sign lepton pair.

3 AAD 15Z based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Used events with  $\ell + \cancel{E}_T + \geq 6j$  ( $\geq 1b$ ) and at least one pair of jets from weak boson decay, primarily designed to select the signature  $b'\bar{b}' \rightarrow WWt\bar{t} \rightarrow WWWWb\bar{b}$ . This is a limit on pair-produced vector-like  $b'$ . The lower mass limit is 640 GeV for a vector-like singlet  $b'$ .

4 Based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. No significant excess over SM expectation is found in the search for pair production or single production of  $b'$  in the events with dilepton from a high  $p_T$   $Z$  and additional jets ( $\geq 1b$ -tag). If instead of  $B(b' \rightarrow Wt) = 1$  an electroweak singlet with  $B(b' \rightarrow Wt) \sim 0.45$  is assumed, the limit reduces to 685 GeV.

5 Based on 5.0 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 13i looked for events with one isolated electron or muon, large  $\cancel{E}_T$ , and at least four jets with large transverse momenta, where one jet is likely to originate from the decay of a bottom quark.

6 Result is based on 1.1 fb<sup>-1</sup> of data. No signal is found for the search of long-lived particles which decay into final states with two electrons or photons, and upper bound on the cross section times branching fraction is obtained for  $2 < c\tau < 7000$  mm; see Fig. 3. 95% CL excluded region of  $b'$  lifetime and mass is shown in Fig. 4.

7 ACOSTA 03 looked for long-lived fourth generation quarks in the data sample of 90 pb<sup>-1</sup> of  $\sqrt{s}=1.8$  TeV  $p\bar{p}$  collisions by using the muon-like penetration and anomalously high ionization energy loss signature. The corresponding lower mass bound for the charge (2/3) $e$  quark ( $t'$ ) is 220 GeV. The  $t'$  bound is higher than the  $b'$  bound because  $t'$  is more likely to produce charged hadrons than  $b'$ . The 95% CL upper bounds for the production cross sections are given in their Fig. 3.

8 AAD 15AR based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Used lepton-plus-jets final state. See Fig. 24 for mass limits in the plane of  $B(b' \rightarrow Wt)$  vs.  $B(b' \rightarrow Hb)$  from  $b'\bar{b}' \rightarrow Hb + X$  searches.

9 AAD 15CN based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Limit on pair-production of chiral  $b'$ -quark. Used events with  $\ell + \cancel{E}_T + \geq 4j$  (non- $b$ -tagged). Limits on a heavy vector-like quark, which decays into  $Wq$ ,  $Zq$ ,  $hq$ , are presented in the plane  $B(Q \rightarrow Wq)$  vs.  $B(Q \rightarrow hq)$  in Fig. 12.

10 Based on 1.04 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. No signal is found for the search of heavy quark pair production that decay into  $W$  and a  $t$  quark in the events with a high  $p_T$  isolated lepton, large  $\cancel{E}_T$ , and at least 6 jets in which one, two or more dijets are from  $W$ .

11 Based on 2.0 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. No  $b' \rightarrow Zb$  invariant mass peak is found in the search of heavy quark pair production that decay into  $Z$  and a  $b$  quark in events with  $Z \rightarrow e^+e^-$  and at least one  $b$ -jet. The lower mass limit is 358 GeV for a vector-like singlet  $b'$  mixing solely with the third SM generation.

12 Based on 1.04 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. No signal is found for the search of heavy quark pair production that decay into  $W$  and a quark in the events with dileptons, large  $\cancel{E}_T$ , and  $\geq 2$  jets.

13 Based on 1.04 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. AAD 12bE looked for events with two isolated like-sign leptons and at least 2 jets, large  $\cancel{E}_T$  and  $H_T > 350$  GeV.

14 Based on 5 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 12bH searched for QCD and EW production of single and pair of degenerate 4<sup>th</sup> generation quarks that decay to  $bW$  or  $tW$ . Absence of signal in events with one lepton, same-sign dileptons or tripletons gives the bound. With a mass difference of 25 GeV/ $c^2$  between  $m_{t'}$  and  $m_{b'}$ , the corresponding limit shifts by about  $\pm 20$  GeV/ $c^2$ .

15 Based on 4.9 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 12x looked for events with tripletons or same-sign dileptons and at least one  $b$  jet.

16 Based on 4.8 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. AALTONEN 11j looked for events with  $\ell + \cancel{E}_T + \geq 5j$  ( $\geq 1b$  or  $c$ ). No signal is observed and the bound  $\sigma(b'\bar{b}')$  < 30 fb for  $m_{b'} > 375$  GeV is found for  $B(b' \rightarrow Wt) = 1$ .

17 Based on 34 pb<sup>-1</sup> of data in  $pp$  collisions at 7 TeV. CHATRCHYAN 11L looked for multi-jet events with tripletons or same-sign dileptons. No excess above the SM background excludes  $m_{b'}$  between 255 and 361 GeV at 95% CL for  $B(b' \rightarrow Wt) = 1$ .

18 Based on 2.7 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV. AALTONEN 10H looked for pair production of heavy quarks which decay into  $tW^-$  or  $tW^+$ , in events with same sign dileptons ( $e$  or  $\mu$ ), several jets and large missing  $E_T$ . The result is obtained for  $b'$  which decays into  $tW^-$ . For the charge 5/3 quark ( $T_{5/3}$ ) which decays into  $tW^+$ ,  $m_{T_{5/3}} > 365$  GeV (95% CL) is found when it has the charge  $-1/3$  partner  $B$  of the same mass.

19 FLACCO 10 result is obtained from AALTONEN 10H result of  $m_{b'} > 338$  GeV, by relaxing the condition  $B(b' \rightarrow Wt) = 100\%$  when  $m_{b'} > m_{t'}$ .

20 Result is based on 1.06 fb<sup>-1</sup> of data. No excess from the SM  $Z$ +jet events is found when  $Z$  decays into  $ee$  or  $\mu\mu$ . The  $m_{b'}$  bound is found by comparing the resulting upper bound on  $\sigma(b'\bar{b}')$  [ $1-(1-B(b' \rightarrow Zb))$ ] and the LO estimate of the  $b'$  pair production cross section shown in Fig. 38 of the article.

21 HUANG 08 reexamined the  $b'$  mass lower bound of 268 GeV obtained in AALTONEN 07c that assumes  $B(b' \rightarrow Zb) = 1$ , which does not hold for  $m_{b'} > 255$  GeV. The lower mass bound is given in the plane of  $\sin^2(\theta_{b'})$  and  $m_{b'}$ .

22 AFFOLDER 00 looked for  $b'$  that decays in to  $b+Z$ . The signal searched for is  $bbZZ$  events where one  $Z$  decays into  $e^+e^-$  or  $\mu^+\mu^-$  and the other  $Z$  decays hadronically. The bound assumes  $B(b' \rightarrow Zb) = 100\%$ . Between 100 GeV and 199 GeV, the 95%CL upper bound on  $\sigma(b' \rightarrow \bar{b}') \times B^2(b' \rightarrow Zb)$  is also given (see their Fig. 2).

23 ABE 98N looked for  $Z \rightarrow e^+e^-$  decays with displaced vertices. Quoted limit assumes  $B(b' \rightarrow Zb) = 1$  and  $c\tau_{b'} = 1$  cm. The limit is lower than  $m_Z + m_b$  ( $\sim 96$  GeV) if  $c\tau > 22$  cm or  $c\tau < 0.009$  cm. See their Fig. 4.

24 ABACHI 97D searched for  $b'$  that decays mainly via FCNC. They obtained 95%CL upper bounds on  $B(b'\bar{b}' \rightarrow \gamma + 3 \text{ jets})$  and  $B(b'\bar{b}' \rightarrow 2\gamma + 2 \text{ jets})$ , which can be interpreted as the lower mass bound  $m_{b'} > m_Z + m_b$ .

25 ABACHI 95F bound on the top-quark also applies to  $b'$  and  $t'$  quarks that decay predominantly into  $W$ . See FROGGATT 97.

26 MUKHOPADHYAYA 93 analyze CDF dilepton data of ABE 92G in terms of a new quark decaying via flavor-changing neutral current. The above limit assumes  $B(b' \rightarrow b\ell^+\ell^-) = 1\%$ . For an exotic quark decaying only via virtual  $Z$  [ $B(b\ell^+\ell^-) = 3\%$ ], the limit is 85 GeV.

27 ABE 92 dilepton analysis limit of >85 GeV at CL=95% also applies to  $b'$  quarks, as discussed in ABE 90B.

28 ABE 90B exclude the region 28–72 GeV.

29 AKESSON 90 searched for events having an electron with  $p_T > 12$  GeV, missing momentum > 15 GeV, and a jet with  $E_T > 10$  GeV,  $|\eta| < 2.2$ , and excluded  $m_{b'}$  between 30 and 69 GeV.

30 For the reduction of the limit due to non-charged-current decay modes, see Fig. 19 of ALBAJAR 90B.

31 ALBAJAR 88 study events at  $E_{cm} = 546$  and 630 GeV with a muon or isolated electron, accompanied by one or more jets and find agreement with Monte Carlo predictions for the production of charm and bottom, without the need for a new quark. The lower mass limit is obtained by using a conservative estimate for the  $b'\bar{b}'$  production cross section and by assuming that it cannot be produced in  $W$  decays. The value quoted here is revised using the full  $O(\alpha_s^3)$  cross section of ALTARELLI 88.

#### $b'(-1/3)$ mass limits from single production in $p\bar{p}$ and $pp$ collisions

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>1390	95	1 KHACHATRY..16i	CMS	$g b \rightarrow b' \rightarrow tW, B(b' \rightarrow tW)=1$
>1430	95	2 KHACHATRY..16i	CMS	$g b \rightarrow b' \rightarrow tW, B(b' \rightarrow tW)=1$
>1530	95	3 KHACHATRY..16i	CMS	$g b \rightarrow b' \rightarrow tW, B(b' \rightarrow tW)=1$
> 693	95	4 ABAZOV	11F D0	$q u \rightarrow q' b' \rightarrow q'(Wu)$ $\tilde{\kappa}_{ub'}=1, B(b' \rightarrow Wu)=1$
> 430	95	4 ABAZOV	11F D0	$q d \rightarrow q b' \rightarrow q(Zd)$ $\tilde{\kappa}_{db'}=\sqrt{2}, B(b' \rightarrow Zd)=1$

1 Based on 19.7 fb<sup>-1</sup> of data in  $pp$  collisions at 8 TeV. Limit on left-handed  $b'$  assuming 100% decay to  $tW$  and using all-hadronic, lepton + jets, and dilepton final states.

2 Based on 19.7 fb<sup>-1</sup> of data in  $pp$  collisions at 8 TeV. Limit on right-handed  $b'$  assuming 100% decay to  $tW$  and using all-hadronic, lepton + jets, and dilepton final states.

3 Based on 19.7 fb<sup>-1</sup> of data in  $pp$  collisions at 8 TeV. Limit on vector-like  $b'$  assuming 100% decay to  $tW$  and using all-hadronic, lepton+ jets, and dilepton final states.

4 Based on 5.4 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. ABAZOV 11F looked for single production of  $b'$  via the  $W$  or  $Z$  coupling to the first generation up or down quarks, respectively. Model independent cross section limits for the single production processes  $p\bar{p} \rightarrow b'q \rightarrow Wuq$ , and  $p\bar{p} \rightarrow b'q \rightarrow Zdq$  are given in Figs. 3 and 4,



See key on page 601

## Quark Particle Listings

 $b'$  (Fourth Generation) Quark,  $t'$  (Fourth Generation) Quark

respectively, and the mass limits are obtained for the model of ATRE 09 with degenerate bi-doublets of vector-like quarks.

MASS LIMITS for  $b'$  (4<sup>th</sup> Generation) Quark or Hadron in  $e^+e^-$  Collisions

Search for hadrons containing a fourth-generation  $-1/3$  quark denoted  $b'$ .

The last column specifies the assumption for the decay mode ( $CC$  denotes the conventional charged-current decay) and the event signature which is looked for.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;46.0</b>	95	<sup>1</sup> DECAMP	90F ALEP	any decay
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 96–103	95	<sup>2</sup> ABDALLAH	07 DLPH	$b' \rightarrow bZ, cW$
		<sup>3</sup> ADRIANI	93G L3	Quarkonium
>44.7	95	ADRIANI	93M L3	$\Gamma(Z)$
>45	95	ABREU	91F DLPH	$\Gamma(Z)$
none 19.4–28.2	95	ABE	90D VNS	Any decay; event shape
>45.0	95	ABREU	90D DLPH	$B(C C) = 1$ ; event shape
>44.5	95	<sup>4</sup> ABREU	90D DLPH	$b' \rightarrow cH^-, H^- \rightarrow \bar{\nu}_s, \tau^- \nu$
>40.5	95	<sup>5</sup> ABREU	90D DLPH	$\Gamma(Z) \rightarrow$ hadrons
>28.3	95	ADACHI	90 TOPZ	$B(FCNC)=100\%$ ; isol. $\gamma$ or 4 jets
>41.4	95	<sup>6</sup> AKRAWY	90B OPAL	Any decay; acoplanarity
>45.2	95	<sup>6</sup> AKRAWY	90B OPAL	$B(C C) = 1$ ; acoplanarity
>46	95	<sup>7</sup> AKRAWY	90J OPAL	$b' \rightarrow \gamma + \text{any}$
>27.5	95	<sup>8</sup> ABE	89E VNS	$B(C C) = 1$ ; $\mu, e$
none 11.4–27.3	95	<sup>9</sup> ABE	89G VNS	$B(b' \rightarrow b\gamma) > 10\%$ ; isolated $\gamma$
>44.7	95	<sup>10</sup> ABRAMS	89C MRK2	$B(C C) = 100\%$ ; isol. track
>42.7	95	<sup>10</sup> ABRAMS	89C MRK2	$B(bg) = 100\%$ ; event shape
>42.0	95	<sup>10</sup> ABRAMS	89C MRK2	Any decay; event shape
>28.4	95	<sup>11,12</sup> ADACHI	89C TOPZ	$B(C C) = 1$ ; $\mu$
>28.8	95	<sup>13</sup> ENO	89 AMY	$B(C C) \gtrsim 90\%$ ; $\mu, e$
>27.2	95	<sup>13,14</sup> ENO	89 AMY	any decay; event shape
>29.0	95	<sup>13</sup> ENO	89 AMY	$B(b' \rightarrow bg) \gtrsim 85\%$ ; event shape
>24.4	95	<sup>15</sup> IGARASHI	88 AMY	$\mu, e$
>23.8	95	<sup>16</sup> SAGAWA	88 AMY	event shape
>22.7	95	<sup>17</sup> ADEVA	86 MRKJ	$\mu$
>21		<sup>18</sup> ALTHOFF	84C TASS	$R$ , event shape
>19		<sup>19</sup> ALTHOFF	84I TASS	Aplanarity

<sup>1</sup> DECAMP 90F looked for isolated charged particles, for isolated photons, and for four-jet final states. The modes  $b' \rightarrow bg$  for  $B(b' \rightarrow bg) > 65\%$   $b' \rightarrow b\gamma$  for  $B(b' \rightarrow b\gamma) > 5\%$  are excluded. Charged Higgs decay were not discussed.

<sup>2</sup> ABDALLAH 07 searched for  $b'$  pair production at  $E_{\text{cm}}=196\text{--}209$  GeV, with  $420\text{ pb}^{-1}$ . No signal leads to the 95% CL upper limits on  $B(b' \rightarrow bZ)$  and  $B(b' \rightarrow cW)$  for  $m_{b'} = 96$  to  $103$  GeV.

<sup>3</sup> ADRIANI 93G search for vector quarkonium states near  $Z$  and give limit on quarkonium- $Z$  mixing parameter  $\delta m^2 < (10\text{--}30)$  GeV<sup>2</sup> (95%CL) for the mass  $88\text{--}94.5$  GeV. Using Richardson potential, a  $1S$  ( $b'\bar{b}'$ ) state is excluded for the mass range  $87.7\text{--}94.7$  GeV. This range depends on the potential choice.

<sup>4</sup> ABREU 90D assumed  $m_{H^-} < m_{b'} - 3$  GeV.

<sup>5</sup> Superseded by ABREU 91F.

<sup>6</sup> AKRAWY 90B search was restricted to data near the  $Z$  peak at  $E_{\text{cm}} = 91.26$  GeV at LEP. The excluded region is between  $23.6$  and  $41.4$  GeV if no  $H^+$  decays exist. For charged Higgs decays the excluded regions are between  $(m_{H^+} + 1.5 \text{ GeV})$  and  $45.5$  GeV.

<sup>7</sup> AKRAWY 90J search for isolated photons in hadronic  $Z$  decay and derive  $B(Z \rightarrow b'\bar{b}')B(b' \rightarrow \gamma X)/B(Z \rightarrow \text{hadrons}) < 2.2 \times 10^{-3}$ . Mass limit assumes  $B(b' \rightarrow \gamma X) > 10\%$ .

<sup>8</sup> ABE 89E search at  $E_{\text{cm}} = 56\text{--}57$  GeV at TRISTAN for multihadron events with a spherical shape (using thrust and acoplanarity) or containing isolated leptons.

<sup>9</sup> ABE 89G search was at  $E_{\text{cm}} = 55\text{--}60.8$  GeV at TRISTAN.

<sup>10</sup> If the photonic decay mode is large ( $B(b' \rightarrow b\gamma) > 25\%$ ), the ABRAMS 89C limit is  $45.4$  GeV. The limit for Higgs decay ( $b' \rightarrow cH^-, H^- \rightarrow \bar{\nu}_s$ ) is  $45.2$  GeV.

<sup>11</sup> ADACHI 89C search was at  $E_{\text{cm}} = 56.5\text{--}60.8$  GeV at TRISTAN using multi-hadron events accompanying muons.

<sup>12</sup> ADACHI 89C also gives limits for any mixture of  $CC$  and  $bg$  decays.

<sup>13</sup> ENO 89 search at  $E_{\text{cm}} = 50\text{--}60.8$  at TRISTAN.

<sup>14</sup> ENO 89 considers arbitrary mixture of the charged current,  $bg$ , and  $b\gamma$  decays.

<sup>15</sup> IGARASHI 88 searches for leptons in low-thrust events and gives  $\Delta R(b') < 0.26$  (95% CL) assuming charged current decay, which translates to  $m_{b'} > 24.4$  GeV.

<sup>16</sup> SAGAWA 88 set limit  $\sigma(\text{top}) < 6.1$  pb at CL=95% for top-flavored hadron production from event shape analyses at  $E_{\text{cm}} = 52$  GeV. By using the quark parton model cross-section formula near threshold, the above limit leads to lower mass bounds of  $23.8$  GeV for charge  $-1/3$  quarks.

<sup>17</sup> ADEVA 86 give 95%CL upper bound on an excess of the normalized cross section,  $\Delta R$ , as a function of the minimum c.m. energy (see their figure 3). Production of a pair of  $1/3$  charge quarks is excluded up to  $E_{\text{cm}} = 45.4$  GeV.

<sup>18</sup> ALTHOFF 84C narrow state search sets limit  $\Gamma(e^+e^-)B(\text{hadrons}) < 2.4$  keV CL = 95% and heavy charge  $1/3$  quark pair production  $m > 21$  GeV, CL = 95%.

<sup>19</sup> ALTHOFF 84I exclude heavy quark pair production for  $7 < m < 19$  GeV ( $1/3$  charge) using aplanarity distributions (CL = 95%).

REFERENCES FOR Searches for (Fourth Generation)  $b'$  Quark

KHACHATRYAN...	16I	JHEP	1601	166	V. Khachatryan et al.	(CMS Collab.)
AAD	15AR	JHEP	1508	105	G. Aad et al.	(ATLAS Collab.)
AAD	15BY	JHEP	1510	150	G. Aad et al.	(ATLAS Collab.)
AAD	15CM	PR	D92	112007	G. Aad et al.	(ATLAS Collab.)
AAD	15Z	PR	D91	112011	G. Aad et al.	(ATLAS Collab.)
AAD	14AZ	JHEP	1411	104	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN	13I	JHEP	1301	154	S. Chatrchyan et al.	(CMS Collab.)
AAD	12AT	PRL	109	032001	G. Aad et al.	(ATLAS Collab.)
AAD	12AU	PRL	109	071801	G. Aad et al.	(ATLAS Collab.)
AAD	12BC	PR	D86	012007	G. Aad et al.	(ATLAS Collab.)
AAD	12BE	JHEP	1204	069	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN	12BH	PR	D86	112003	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12X	JHEP	1205	123	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN	11J	PRL	106	141803	T. Aaltonen et al.	(CDF Collab.)
ABAZOV	11F	PRL	106	081801	V.M. Abazov et al.	(DO Collab.)
CHATRCHYAN	11L	PL	B701	204	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN	10H	PRL	104	091801	T. Aaltonen et al.	(CDF Collab.)
FLACCO	10	PRL	105	111801	C.J. Flacco et al.	(UCI, HAIF)
ATRE	09	PR	D79	054018	A. Atre et al.	
ABAZOV	08X	PRL	101	111802	V.M. Abazov et al.	(DO Collab.)
HUANG	08	PR	D77	037302	P.Q. Hung, M. Sher	(UVA, WILL)
AALTONEN	07C	PR	D76	072006	T. Aaltonen et al.	(CDF Collab.)
ABDALLAH	07	EPJ	C50	507	J. Abdallah et al.	(DELPHI Collab.)
ACOSTA	03	PRL	90	131801	D. Acosta et al.	(CDF Collab.)
AFFOLDER	00	PRL	84	835	A. Affolder et al.	(CDF Collab.)
ABE	98N	PR	D58	051102	F. Abe et al.	(CDF Collab.)
ABACHI	97D	PRL	78	3818	S. Abachi et al.	(DO Collab.)
FROGGATT	97	ZPHY	C73	333	C.D. Froggatt, D.J. Smith, H.B. Nielsen	(GLAS+)
ABACHI	95F	PR	D52	4877	S. Abachi et al.	(DO Collab.)
ADRIANI	93G	PL	B313	326	O. Adriani et al.	(L3 Collab.)
ADRIANI	93M	PRPL	236	1	O. Adriani et al.	(L3 Collab.)
MUKHOPADH...	93	PR	D48	2105	B. Mukhopadhyaya, D.P. Roy	(TATA)
ABE	92	PRL	68	447	F. Abe et al.	(CDF Collab.)
Also					F. Abe et al.	(CDF Collab.)
ABE	92G	PR	D45	3921	F. Abe et al.	(CDF Collab.)
ABREU	91F	NP	B367	511	P. Abreu et al.	(DELPHI Collab.)
ABE	90B	PRL	64	147	F. Abe et al.	(CDF Collab.)
ABE	90D	PL	B234	382	K. Abe et al.	(VENUS Collab.)
ABREU	90D	PL	B242	536	P. Abreu et al.	(DELPHI Collab.)
ADACHI	90	PL	B234	197	I. Adachi et al.	(TOPAZ Collab.)
AKESSON	90	ZPHY	C46	179	T. Akesson et al.	(UA2 Collab.)
AKRAWY	90B	PL	B236	364	M.Z. Akrawy et al.	(OPAL Collab.)
AKRAWY	90J	PL	B246	285	M.Z. Akrawy et al.	(OPAL Collab.)
ALBAJAR	90B	ZPHY	C48	1	C. Albajar et al.	(UA1 Collab.)
DECAMP	90J	PL	B236	511	D. Decamp et al.	(ALEPH Collab.)
ABE	89E	PR	D39	3524	K. Abe et al.	(VENUS Collab.)
ABE	89G	PRL	63	1776	K. Abe et al.	(VENUS Collab.)
ABRAMS	89C	PRL	63	2447	G.S. Abrams et al.	(Mark II Collab.)
ADACHI	89C	PL	B229	427	I. Adachi et al.	(TOPAZ Collab.)
ENO	89	PRL	63	1910	S. Eno et al.	(AMY Collab.)
ALBAJAR	88	ZPHY	C37	505	C. Albajar et al.	(UA1 Collab.)
ALTARELLI	88	NP	B308	724	G. Altarelli et al.	(CERN, ROMA, ETH)
IGARASHI	88	PL	60	2359	S. Igarashi et al.	(AMY Collab.)
SAGAWA	88	PRL	60	93	H. Sagawa et al.	(AMY Collab.)
ADEVA	86	PR	D34	681	B. Adeva et al.	(Mark-I Collab.)
ALTHOFF	84C	PL	138B	441	M. Althoff et al.	(TASSO Collab.)
ALTHOFF	84I	ZPHY	C22	307	M. Althoff et al.	(TASSO Collab.)

 $t'$  (4<sup>th</sup> Generation) Quark, Searches for $t'(2/3)$ -quark/hadron mass limits in  $p\bar{p}$  and  $pp$  collisions

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>770	95	<sup>1</sup> AAD	15AR ATLS	$B(t' \rightarrow Wb) = 1$
>590	95	<sup>2</sup> AAD	15BY ATLS	$Wb, Zt, ht$ modes
>745	95	<sup>3</sup> KHACHATRYAN...	15AI CMS	$B(t' \rightarrow ht) = 1$
>735	95	<sup>4</sup> AAD	14AZ ATLS	
>700	95	<sup>5</sup> CHATRCHYAN14A	CMS	$B(t' \rightarrow Wb) = 1$
>706	95	<sup>5</sup> CHATRCHYAN14A	CMS	$B(t' \rightarrow Zt) = 1$
>782	95	<sup>5</sup> CHATRCHYAN14A	CMS	$B(t' \rightarrow ht) = 1$
>350	95	<sup>6</sup> AAD	12BC ATLS	$B(t' \rightarrow Wq)=1$ ( $q=d,s,b$ )
>420	95	<sup>7</sup> AAD	12C ATLS	$t' \rightarrow X t$ ( $m_X < 140$ GeV)
>685	95	<sup>8</sup> CHATRCHYAN12BH	CMS	$m_{b'} = m_{t'}$
>557	95	<sup>9</sup> CHATRCHYAN12P	CMS	$t'\bar{t}' \rightarrow W^+bW^-\bar{b} \rightarrow b\bar{t}'\bar{\nu}\bar{b}\ell^+\bar{\nu}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

>656	95	<sup>10</sup> AAD	13F ATLS	$B(t' \rightarrow Wb) = 1$
>625	95	<sup>11</sup> CHATRCHYAN13I	CMS	$B(t' \rightarrow Zt) = 1$
>404	95	<sup>12</sup> AAD	12AR ATLS	$B(t' \rightarrow Wb) = 1$
>570	95	<sup>13</sup> CHATRCHYAN12BC	CMS	$t'\bar{t}' \rightarrow W^+bW^-\bar{b}$
>400	95	<sup>14</sup> AALTONEN	11AH CDF	$t' \rightarrow X t$ ( $m_X < 70$ GeV)
>358	95	<sup>15</sup> AALTONEN	11AL CDF	$t' \rightarrow Wb$
>340	95	<sup>15</sup> AALTONEN	11AL CDF	$t' \rightarrow Wq$ ( $q=d,s,b$ )
>360	95	<sup>16</sup> AALTONEN	11O CDF	$t' \rightarrow X t$ ( $m_X < 100$ GeV)
>285	95	<sup>17</sup> ABAZOV	11Q D0	$t' \rightarrow Wq$ ( $q=d,s,b$ )
>256	95	<sup>18,19</sup> AALTONEN	08H CDF	$t' \rightarrow Wq$

<sup>1</sup> AAD 15AR based on  $20.3\text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 8$  TeV. Used lepton-plus-jets final state. See Fig. 20 for mass limits in the plane of  $B(t' \rightarrow Ht)$  vs.  $B(t' \rightarrow Wb)$  from a combination of  $t'\bar{t}' \rightarrow Wb + X$  and  $t'\bar{t}' \rightarrow Ht + X$  searches. Any branching ratio scenario is excluded for mass below  $715$  GeV.

<sup>2</sup> AAD 15BY based on  $20.3\text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 8$  TeV. Limit on pair-produced vector-like  $t'$  assuming the branching fractions to  $W, Z$ , and  $h$  modes of the singlet lepton pair. Used events containing  $\geq 2\ell + \cancel{E}_T + \geq 2j$  ( $\geq 1b$ ) and including a same-sign lepton pair.

<sup>3</sup> KHACHATRYAN 15AI based on  $19.7\text{ fb}^{-1}$  of  $pp$  data at  $\sqrt{s} = 8$  TeV. The search exploits all-hadronic final states by tagging boosted Higgs boson using jet substructure and  $b$ -tagging.

# Quark Particle Listings

## $t'$ (Fourth Generation) Quark, Free Quark Searches

- <sup>4</sup> Based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. No significant excess over SM expectation is found in the search for pair production or single production of  $t'$  in the events with dilepton from a high  $p_T$   $Z$  and additional jets ( $\geq 1$   $b$ -tag). If instead of  $B(b' \rightarrow Wt) = 1$  an electroweak singlet with  $B(b' \rightarrow Wt) \sim 0.45$  is assumed, the limit reduces to 685 GeV.
- <sup>5</sup> Based on 19.5 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. The  $t'$  quark is pair produced and is assumed to decay into three different final states of  $bW$ ,  $tZ$ , and  $th$ . The search is carried out using events with at least one isolated lepton.
- <sup>6</sup> Based on 1.04 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. No signal is found for the search of heavy quark pair production that decay into  $W$  and a quark in the events with dileptons, large  $E_T$ , and  $\geq 2$  jets.
- <sup>7</sup> Based on 1.04 fb<sup>-1</sup> of data in  $pp$  collisions at 7 TeV. AAD 12c looked for  $t'\bar{t}'$  production followed by  $t'$  decaying into a top quark and  $X$ , an invisible particle, in a final state with an isolated high- $p_T$  lepton, four or more jets, and a large missing transverse energy. No excess over the SM  $t\bar{t}$  production gives the upper limit on  $t'\bar{t}'$  production cross section as a function of  $m_{t'}$  and  $m_X$ . The result is obtained for  $B(t' \rightarrow Wt) = 1$ .
- <sup>8</sup> Based on 5 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 12BH searched for QCD and EW production of single and pair of degenerate 4<sup>th</sup> generation quarks that decay to  $Wb$  or  $Wt$ . Absence of signal in events with one lepton, same-sign dileptons or tripletons gives the bound. With a mass difference of 25 GeV/c<sup>2</sup> between  $m_{t'}$  and  $m_{b'}$ , the corresponding limit shifts by about  $\pm 20$  GeV/c<sup>2</sup>.
- <sup>9</sup> Based on 5.0 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 12P looked for  $t'\bar{t}'$  production events with two isolated high  $p_T$  leptons, large  $E_T$ , and 2 high  $p_T$  jets with  $b$ -tag. The absence of signal above the SM background gives the limit for  $B(t' \rightarrow Wb) = 1$ .
- <sup>10</sup> Based on 4.7 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. No signal is found for the search of heavy quark pair production that decay into  $W$  and a  $b$  quark in the events with a high  $p_T$  isolated lepton, large  $E_T$  and at least 3 jets ( $\geq 1$   $b$ -tag). Vector-like quark of charge 2/3 with  $400 < m_{t'} < 550$  GeV and  $B(t' \rightarrow Wb) > 0.63$  is excluded at 95% CL.
- <sup>11</sup> Based on 5.0 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 13I looked for events with one isolated electron or muon, large  $E_T$ , and at least four jets with large transverse momenta, where one jet is likely to originate from the decay of a bottom quark.
- <sup>12</sup> Based on 1.04 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. No signal is found in the search for pair produced heavy quarks that decay into  $W$  boson and a  $b$  quark in the events with a high  $p_T$  isolated lepton, large  $E_T$  and at least 3 jets ( $\geq 1$   $b$ -tag).
- <sup>13</sup> Based on 5.0 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 7$  TeV. CHATRCHYAN 12Bc looked for  $t'\bar{t}'$  production events with a single isolated high  $p_T$  lepton, large  $E_T$  and at least 4 high  $p_T$  jets with a  $b$ -tag. The absence of signal above the SM background gives the limit for  $B(t' \rightarrow Wb) = 1$ .
- <sup>14</sup> Based on 5.7 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. AALTONEN 11AH looked for  $t'\bar{t}'$  production followed by  $t'$  decaying into a top quark and  $X$ , an invisible particle, in the all hadronic decay mode of  $t\bar{t}$ . No excess over the SM  $t\bar{t}$  production gives the upper limit on  $t'\bar{t}'$  production cross section as a function of  $m_{t'}$  and  $m_X$ . The result is obtained for  $B(t' \rightarrow Xt) = 1$ .
- <sup>15</sup> Based on 5.6 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. AALTONEN 11AL looked for  $\ell + \geq 4j$  events and set upper limits on  $\sigma(t'\bar{t}')$  as functions of  $m_{t'}$ .
- <sup>16</sup> Based on 4.8 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. AALTONEN 11I looked for  $t'\bar{t}'$  production signal when  $t'$  decays into a top quark and  $X$ , an invisible particle, in  $\ell + E_T + \text{jets}$  channel. No excess over the SM  $t\bar{t}$  production gives the upper limit on  $t'\bar{t}'$  production cross section as a function of  $m_{t'}$  and  $m_X$ . The result is obtained for  $B(t' \rightarrow Xt) = 1$ .
- <sup>17</sup> Based on 5.3 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. ABZOV 11Q looked for  $\ell + E_T + \geq 4j$  events and set upper limits on  $\sigma(t'\bar{t}')$  as functions of  $m_{t'}$ .
- <sup>18</sup> Searches for pair production of a new heavy top-like quark  $t'$  decaying to a  $W$  boson and another quark by fitting the observed spectrum of total transverse energy and reconstructed  $t'$  mass in the lepton + jets events.
- <sup>19</sup> HUANG 08 reexamined the  $t'$  mass lower bound of 256 GeV obtained in AALTONEN 08H that assumes  $B(b' \rightarrow qZ) = 1$  for  $q = u, c$  which does not hold when  $m_{b'} < m_{t'} - m_W$  or the mixing  $\sin^2(\theta_{bt'})$  is so tiny that the decay occurs outside of the vertex detector. Fig. 1 gives that lower bound on  $m_{t'}$  in the plane of  $\sin^2(\theta_{bt'})$  and  $m_{b'}$ .

### $t'(5/3)$ -quark/hadron mass limits in $p\bar{p}$ and $pp$ collisions

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>750	95	<sup>1</sup> AAD 15BY ATLS	$t'(5/3) \rightarrow tW^+$	
>840	95	<sup>2</sup> AAD 15Z ATLS	$t'(5/3) \rightarrow tW^+$	
>800	95	<sup>3</sup> CHATRCHYAN 14T	CMS $t'(5/3) \rightarrow tW^+$	

<sup>1</sup> AAD 15BY based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Limit on  $t'(5/3)$  in pair and single production assuming its coupling to  $Wt$  is equal to one. Used events containing  $\geq 2\ell + E_T + \geq 2j$  ( $\geq 1$   $b$ ) and including a same-sign lepton pair.

<sup>2</sup> AAD 15Z based on 20.3 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Used events with  $\ell + E_T + \geq 6j$  ( $\geq 1$   $b$ ) and at least one pair of jets from weak boson decay, sensitive to the final state  $b\bar{b}W^+W^-W^+W^-$ .

<sup>3</sup> Based on 19.5 fb<sup>-1</sup> of  $pp$  data at  $\sqrt{s} = 8$  TeV. Non-observation of anomaly in  $H_T$  distribution in the same sign dilepton events leads to the limit when pair produced  $t'(5/3)$  quark decays exclusively into  $t$  and  $W^+$ , resulting in the final state with  $b\bar{b}W^+W^-W^+W^-$ .

### $t'(2/3)$ mass limits from single production in $p\bar{p}$ and $pp$ collisions

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>403	95	<sup>1</sup> ABZOV	11F D0	$q\bar{d} \rightarrow q't' \rightarrow q'(Wd)$ $\bar{\kappa}_{d't'}=1, B(t' \rightarrow Wd)=1$
>551	95	<sup>1</sup> ABZOV	11F D0	$qu \rightarrow q't' \rightarrow q(Zu)$ $\bar{\kappa}_{u't'}=\sqrt{2}, B(t' \rightarrow Zu)=1$

<sup>1</sup> Based on 5.4 fb<sup>-1</sup> of data in  $p\bar{p}$  collisions at 1.96 TeV. ABZOV 11F looked for single production of  $t'$  via the  $Z$  or  $E$  coupling to the first generation up or down quarks, respectively. Model independent cross section limits for the single production processes  $p\bar{p} \rightarrow t'q \rightarrow (Wd)q$ , and  $p\bar{p} \rightarrow t'q \rightarrow (Zd)q$  are given in Figs. 3 and 4, respectively, and the mass limits are obtained for the model of ATRE 09 with degenerate bi-doublets of vector-like quarks.

### REFERENCES FOR Searches for (Fourth Generation) $t'$ Quark

AAD 15AR JHEP 1508 105	G. Aad et al.	(ATLAS Collab.)
AAD 15BY JHEP 1510 150	G. Aad et al.	(ATLAS Collab.)
AAD 15Z PR D91 112011	G. Aad et al.	(ATLAS Collab.)
KHACHATRYAN 15AI JHEP 1506 080	V. Khachatryan et al.	(CMS Collab.)
AAD 14AZ JHEP 1411 104	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN 14A PL B729 149	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 14T PRL 112 171801	S. Chatrchyan et al.	(CMS Collab.)
AAD 13F PL B718 1284	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN 13I JHEP 1301 154	S. Chatrchyan et al.	(CMS Collab.)
AAD 12AR PRL 108 261802	G. Aad et al.	(ATLAS Collab.)
AAD 12BC PR D86 012007	G. Aad et al.	(ATLAS Collab.)
AAD 12C PRL 108 041805	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN 12Bc PL B718 307	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12BH PR D86 112003	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 12P PL B716 103	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN 11AH PRL 107 191803	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 11AL PRL 107 261801	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 11I PRL 106 191801	T. Aaltonen et al.	(CDF Collab.)
ABZOV 11F PRL 106 081801	V.M. Abazov et al.	(DO Collab.)
ABZOV 11Q PRL 107 082001	V.M. Abazov et al.	(DO Collab.)
ATRE 09 PR D79 054018	A. Atre et al.	(CDF Collab.)
AALTONEN 08H PRL 100 161803	T. Aaltonen et al.	(CDF Collab.)
HUANG 08 PR D77 037302	P.Q. Hung, M. Sher	(UVA, WILL)

## Free Quark Searches

### FREE QUARK SEARCHES

The basis for much of the theory of particle scattering and hadron spectroscopy is the construction of the hadrons from a set of fractionally charged constituents (quarks). A central but unproven hypothesis of this theory, Quantum Chromodynamics, is that quarks cannot be observed as free particles but are confined to mesons and baryons.

Experiments show that it is at best difficult to “unglue” quarks. Accelerator searches at increasing energies have produced no evidence for free quarks, while only a few cosmic-ray and matter searches have produced uncorroborated events.

This compilation is only a guide to the literature, since the quoted experimental limits are often only indicative. Reviews can be found in Refs. 1–4.

### References

- M.L. Perl, E.R. Lee, and D. Lomba, Mod. Phys. Lett. **A19**, 2595 (2004).
- P.F. Smith, Ann. Rev. Nucl. and Part. Sci. **39**, 73 (1989).
- L. Lyons, Phys. Reports **129**, 225 (1985).
- M. Marinelli and G. Morpurgo, Phys. Reports **85**, 161 (1982).

### Quark Production Cross Section — Accelerator Searches

X-SECT (cm <sup>2</sup> )	CHG (e/3)	MASS (GeV)	ENERGY (GeV)	BEAM	EVTS	DOCUMENT ID	TECN
<1.7-2.3E-39	±2	100-600	7000	pp	0	<sup>1</sup> CHATRCHYAN 13AR	CMS
<14-5.4E-39	±1	100-600	7000	pp	0	<sup>1</sup> CHATRCHYAN 13AR	CMS
<1.3E-36	±2	45-84	130-172	e <sup>+</sup> e <sup>-</sup>	0	ABREU 97D	DLPH
<2.E-35	+2	250	1800	p $\overline{p}$	0	<sup>2</sup> ABE 92J	CDF
<1.E-35	+4	250	1800	p $\overline{p}$	0	<sup>2</sup> ABE 92J	CDF
<3.8E-28			14.5A	<sup>28</sup> Si-Pb	0	<sup>3</sup> HE 91	PLAS
<3.2E-28			14.5A	<sup>28</sup> Si-Cu	0	<sup>3</sup> HE 91	PLAS
<1.E-40	±1,2	<10		$\rho,\nu,\overline{\nu}$	0	BERGSM 84B	CHRM
<1.E-36	±1,2	<9	200	$\mu$	0	AUBERT 83C	SPEC
<2E-10	±2,4	1-3	200	p	0	<sup>4</sup> BUSSIERE 80	CNTR
<5.E-38	+1,2	>5	300	p	0	<sup>5,6</sup> STEVENSON 79	CNTR
<1.E-33	±1	<20	52	pp	0	BASILE 78	SPEC
<9.E-39	±1,2	<6	400	p	0	<sup>5</sup> ANTREASANY 77	SPEC
<8.E-35	+1,2	<20	52	pp	0	<sup>7</sup> FABJAN 75	CNTR
<5.E-38	-1,2	4-9	200	p	0	NASH 74	CNTR



# Quark Particle Listings

## Free Quark Searches

<8.E-11	+1		8	ASHTON	73	CNTR
<2.E-8	+1,2			HICKS	73B	CNTR
<5.E-10	+4	2.8 *		BEAUCHAMP	72	CNTR
<1.E-10	+1,2		8	BOHM	72B	CNTR
<1.E-10	+1,2	2.8 *		COX	72	ELEC
<3.E-10	+2			CROUCH	72	CNTR
<3.E-8		7	7	DARDO	72	CNTR
<4.E-9	+1		8	EVANS	72	CC
<2.E-9		>10	7	TONWAR	72	CNTR
<2.E-10	+1	2.8 *		CHIN	71	CNTR
<3.E-10	+1,2		8	CLARK	71B	CC
<1.E-10	+1,2		8	HAZEN	71	CC
<5.E-10	+1,2	3.5 *		BOSIA	70	CNTR
	+1,2	<6.5	8	CHU	70	HLBC
<2.E-9	+1			FAISSNER	70B	CNTR
<2.E-10	+1,2	0.8 *		KRIDER	70	CNTR
<5.E-11	+2			CAIRNS	69	CC
<8.E-10	+1,2	<10		FUKUSHIMA	69	CNTR
	+2		8,10	MCCUSKER	69	CC
<1.E-10		>5	1,7,3,6	BJORNBOE	68	CNTR
<1.E-8	±1,2,4	6.3,2 *	5	BRIATORE	68	CNTR
<3.E-8		>2		FRANZINI	68	CNTR
<9.E-11	±1,2			GARMIRE	68	CNTR
<4.E-10	±1			HANAYAMA	68	CNTR
<3.E-8		>15		KASHA	68	OSP K
<2.E-10	+2			KASHA	68B	CNTR
<2.E-10	+4			KASHA	68C	CNTR
<2.E-10	+2	6		BARTON	67	CNTR
<2.E-7	+4	0.008,0.5 *		BUHLER	67	CNTR
<5.E-10	1,2	0.008,0.5 *		BUHLER	67B	CNTR
<4.E-10	+1,2			GOMEZ	67	CNTR
<2.E-9	+2			KASHA	67	CNTR
<2.E-10	+2	220		BARTON	66	CNTR
<2.E-9	+1,2	0.5 *		BUHLER	66	CNTR
<3.E-9	+1,2			KASHA	66	CNTR
<2.E-9	+1,2			LAMB	66	CNTR
<2.E-8	+1,2	>7	2.8 *	DELISE	65	CNTR
<5.E-8	+2	>2.5	0.5 *	MASSAM	65	CNTR
<2.E-8	+1		2.5 *	BOWEN	64	CNTR
<2.E-7	+1		0.8	SUNYAR	64	CNTR

- 1 See AGNESE 15 Fig.6 for limits on vertical density as function of charge extending to  $|q|/e < 1/10$ .
- 2 AMBROSIO 00C limit is below  $11 \times 10^{-15}$  for  $0.25 < q/e < 0.5$ , and is changing rapidly near  $q/e=2/3$ , where it is  $2 \times 10^{-14}$ .
- 3 Distribution in celestial sphere was described as anisotropic.
- 4 With telescope axis at zenith angle  $40^\circ$  to the south.
- 5 Leptonik quarks.
- 6 Lifetime  $> 10^{-8}$  s; charge  $\pm 0.70, 0.68, 0.42$ ; and mass  $> 4.4, 4.8$ , and  $20$  GeV, respectively.
- 7 Time delayed air shower search.
- 8 Prompt air shower search.
- 9 Also  $e/4$  and  $e/6$  charges.
- 10 No events in subsequent experiments.

### Quark Density — Matter Searches

QUARKS/ NUCLEON	CHG (e/3)	MASS (GeV)	MATERIAL/METHOD	EVTS	DOCUMENT ID
<1.17E-22			silicone oil drops	0	1 LEE 02
<4.71E-22			silicone oil drops	1	2 HALYO 00
<4.7E-21	±1,2		silicone oil drops	0	MAR 96
<8.E-22	+2		Si/infrared photoionization	0	PERERA 93
<5.E-27	±1,2		sea water/levitation	0	HOMER 92
<4.E-20	±1,2		meteorites/mag. levitation	0	JONES 89
<1.E-19	±1,2		various/spectrometer	0	MILNER 87
<5.E-22	±1,2		W/levitation	0	SMITH 87
<3.E-20	+1,2		org liq/droplet tower	0	VANPOLEN 87
<6.E-20	-1,2		org liq/droplet tower	0	VANPOLEN 87
<3.E-21	±1		Hg drops-untreated	0	SAVAGE 86
<3.E-22	±1,2		levitated niobium	0	SMITH 86
<2.E-26	±1,2		<sup>4</sup> He/levitation	0	SMITH 86B
<2.E-20	>±1	0.2-250	niobium+tungs/ion	0	MILNER 85
<1.E-21	±1		levitated niobium	0	SMITH 85
	+1,2	<100	niobium/mass spec	0	KUTSCHERA 84
<5.E-22			levitated steel	0	MARINELLI 84
<9.E-20	± <13		water/oil drop	0	JOYCE 83
<2.E-21	>  ± 1/2		levitated steel	0	LIEBOWITZ 83
<1.E-19	±1,2		photo ion spec	0	VANDESTEEG 83
<2.E-20			mercury/oil drop	0	3 HODGES 81
1.E-20	+1		levitated niobium	4	4 LARUE 81
1.E-20	-1		levitated niobium	4	4 LARUE 81
<1.E-21			levitated steel	0	MARINELLI 80B
<3.E-16			helium/mass spec	0	BOYD 79
1.E-20	+1		levitated niobium	2	4 LARUE 79
<4.E-28			earth+/ion beam	0	OGOROD... 79
<5.E-15	+1		tungs./mass spec	0	BOYD 78
<5.E-16	+3	<1.7	hydrogen/mass spec	0	BOYD 78B

<1.E-21	±2,4	water/ion beam	0	LUND	78
<6.E-15	>1/2	levitated tungsten	0	PUTT	78
<1.E-22		metals/mass spec	0	SCHIFFER	78
<5.E-15		levitated tungsten ox	0	BLAND	77
<3.E-21		levitated iron	0	GALLINARO	77
2.E-21	-1	levitated niobium	1	4 LARUE	77
4.E-21	+1	levitated niobium	2	4 LARUE	77
<1.E-13	+3	hydrogen/mass spec	0	MULLER	77
<5.E-27		water+/ion beam	0	OGOROD...	77
<1.E-21		lunar+/ion spec	0	STEVENS	76
<1.E-15	+1	oxygen+/ion spec	0	ELBERT	70
<5.E-19		levitated graphite	0	MORPURGO	70
<5.E-23		water+/atom beam	0	COOK	69
<1.E-17	±1,2	levitated graphite	0	BRAGINSK	68
<1.E-17		water+/uv spec	0	RANK	68
<3.E-19	±1	levitated iron	0	STOVER	67
<1.E-10		sun/uv spec	0	5 BENNETT	66
<1.E-17	+1,2	meteorites+/ion beam	0	CHUPKA	66
<1.E-16	±1	levitated graphite	0	GALLINARO	66
<1.E-22		argon/electrometer	0	HILLAS	59
	-2	levitated oil	0	MILLIKAN	10

- 1 95% CL limit for fractional charge particles with  $0.18e \leq |Q_{residual}| \leq 0.82e$  in total of 70.1 mg of silicone oil.
- 2 95% CL limit for particles with fractional charge  $|Q_{residual}| > 0.16e$  in total of 17.4 mg of silicone oil.
- 3 Also set limits for  $Q = \pm e/6$ .
- 4 Note that in PHILLIPS 88 these authors report a subtle magnetic effect which could account for the apparent fractional charges.
- 5 Limit inferred by JONES 77B.

### REFERENCES FOR Free Quark Searches

AGNESE	15	PRL 114 111302	R. Agnese <i>et al.</i>	(CDMS Collab.)
CHATRCHYAN	13AR	PR D87 092008	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
LEE	02	PR D66 012002	I.T. Lee <i>et al.</i>	
AMBROSIO	00C	PR D62 052003	M. Ambrosio <i>et al.</i>	(MACRO Collab.)
HALYO	00	PRL 84 2576	V. Haljo <i>et al.</i>	
ABREU	97D	PL B396 315	P. Abreu <i>et al.</i>	(DELPHI Collab.)
HUENRUP	96	PR C53 358	G. Huenruep <i>et al.</i>	(SIEG)
MAR	96	PR D53 6017	N.M. Mar <i>et al.</i>	(SLAC, SCHAF, LANL, UCI)
AKERS	95R	ZPHY C67 203	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	93C	PL B303 198	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
CECCHINI	93	ASP 1 369	S. Cecchini <i>et al.</i>	
PERERA	93	PRL 70 1053	A.G.U. Perera <i>et al.</i>	(PITT)
ABE	92J	PR D46 R1889	F. Abe <i>et al.</i>	(CDF Collab.)
GHOSH	92	NC 105A 99	D. Ghosh <i>et al.</i>	(JADA, BANGB)
HOMER	92	ZPHY C55 549	G.J. Homer <i>et al.</i>	(RAL, SHMP, LOQM)
BASILE	91	NC 104A 405	M. Basile <i>et al.</i>	(BGNA, INFN, CERN, PLRM+)
HE	91	PR C44 1672	Y.B. He, P.B. Price	(UCB)
MATIS	91	NP A525 513c	H.S. Matis <i>et al.</i>	(LBL, SFSU, UCI+)
MORI	91	PR D43 2843	M. Mori <i>et al.</i>	(KamioKande II Collab.)
ADACHI	90C	PL B244 352	I. Adachi <i>et al.</i>	(TOPAZ Collab.)
BOWCOCK	89B	PR D40 263	T.J.V. Bowcock <i>et al.</i>	(CLEO Collab.)
CALLOWAY	89	PL B232 549	D. Calloway <i>et al.</i>	(SFSU, UCI, LBL+)
JONES	89	ZPHY C43 349	W.G. Jones <i>et al.</i>	(LOIC, RAL)
MATIS	89	PR D39 1851	H.S. Matis <i>et al.</i>	(LBL, SFSU, UCI+)
NAKAMURA	89	PR D39 1261	T.T. Nakamura <i>et al.</i>	(KYOT, TMTC)
ALLASIA	88	PR D37 219	D. Allasia <i>et al.</i>	(WA25 Collab.)
HOFFMANN	88	PL B200 583	A. Hofmann <i>et al.</i>	(SIEG, USF)
PHILLIPS	88	NIM A264 125	J.D. Phillips, W.M. Fairbank, J. Navarro	(STAN)
WADA	88	NC 11C 229	T. Wada, Y. Yamashita, I. Yamamoto	(OKAY)
GERBIER	87	PRL 59 2535	G. Gerbier <i>et al.</i>	(UCB, CERN)
LYONS	87	ZPHY C56 363	L. Lyons <i>et al.</i>	(OXF, RAL, LOIC)
MILNER	87	PR D36 37	R.E. Milner <i>et al.</i>	(CIT)
SHAW	87	PR D36 3533	G.L. Shaw <i>et al.</i>	(UCI, LBL, LANL, SFSU)
SMITH	87	PL B197 447	P.F. Smith <i>et al.</i>	(RAL, LOIC)
VANPOLEN	87	PR D36 1983	J. van Polen, R.T. Hagstrom, G. Hirsch	(ANL+)
ABACHI	86C	PR D33 2733	S. Abachi <i>et al.</i>	(UCLA, LBL, UCD)
SAVAGE	86	PL 167B 481	M.L. Savage <i>et al.</i>	(SFSU)
SMITH	86	PL B171 129	P.F. Smith <i>et al.</i>	(RAL, LOIC)
SMITH	86B	PL B181 407	P.F. Smith <i>et al.</i>	(RAL, LOIC)
WADA	86	NC 9C 358	T. Wada	(OKAY)
ALBRECHT	85G	PL 156B 134	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BANNER	85	PL 156B 129	M. Banner <i>et al.</i>	(UA2 Collab.)
MILNER	85	PRL 54 1472	R.E. Milner <i>et al.</i>	(CIT)
SMITH	85	PL 153B 188	P.F. Smith <i>et al.</i>	(RAL, LOIC)
AIHARA	84	PRL 52 168	H. Aihara <i>et al.</i>	(TPC Collab.)
AIHARA	84B	PRL 52 2332	H. Aihara <i>et al.</i>	(TPC Collab.)
BARWICK	84	PR D30 691	S.W. Barwick, J.A. Musser, J.D. Stevenson	(UCB)
BERGSMAN	84B	ZPHY C24 217	F. Bergsma <i>et al.</i>	(CHARM Collab.)
BONDAR	84	JETPL 40 1265	A.E. Bondar <i>et al.</i>	(NOVO)
		Translated from ZETFP 40 440.		
GURYN	84	PL 139B 313	W. Guryin <i>et al.</i>	(FRAS, LBL, NWES, STAN+)
KAWAGOE	84B	LNC 44 604	K. Kawagoe <i>et al.</i>	(TOKY)
KUTSCHERA	84	PR D29 791	W. Kutschera <i>et al.</i>	(ANL, FNAL)
MARINELLI	84	PL 137B 439	M. Marinelli, G. Morpurgo	(GENO)
WADA	84B	LNC 40 329	T. Wada, Y. Yamashita, I. Yamamoto	(OKAY)
AUBERT	83C	PL 133B 461	J.J. Aubert <i>et al.</i>	(EMC Collab.)
BANNER	83	PL 121B 187	M. Banner <i>et al.</i>	(UA2 Collab.)
JOYCE	83	PRL 51 731	D.C. Joyce <i>et al.</i>	(SFSU)
LIEBOWITZ	83	PRL 50 1640	D. Liebowitz, M. Binder, K.O.H. Ziock	(UVA)
LINDGREN	83	PRL 51 1621	M.A. Lindgren <i>et al.</i>	(SFSU, UCR, UCI+)
MASHIMO	83	PL 128B 327	T. Mashimo <i>et al.</i>	(ICEPP)
PRICE	83	PRL 50 566	P.B. Price <i>et al.</i>	(UCB)
VANDESTEEG	83	PRL 50 1234	M.J.H. van de Steeg, H.W.H.M. Jongbloets, P. Wyder	(UCL)
MARINI	82B	PR D26 1777	A. Marini <i>et al.</i>	(FRAS, LBL, NWES, STAN+)
MARINI	82B	PRL 48 1649	A. Marini <i>et al.</i>	(FRAS, LBL, NWES, STAN+)
MASHIMO	82	JPSJ 51 3067	T. Mashimo, K. Kawagoe, M. Koshiba	(INUS)
NAPOLITANO	82	PR D25 2837	J. Napolitano <i>et al.</i>	(STAN, FRAS, LBL+)
ROSS	82	PL 118B 199	M.C. Ross <i>et al.</i>	(FRAS, LBL, NWES, STAN+)
HODGES	81	PRL 47 1651	L.C. Hodges <i>et al.</i>	(UCR, SFSU)
LARUE	81	PRL 46 967	G.S. Larue, J.D. Phillips, W.M. Fairbank	(STAN)
WEISS	81	PL 101B 439	J.M. Weiss <i>et al.</i>	(SLAC, LBL, UCB)
BARTLE	80	ZPHY C6 295	W. Bartel <i>et al.</i>	(JADE Collab.)
BASILE	80	LNC 29 251	M. Basile <i>et al.</i>	(BGNA, CERN, FRAS, ROMA+)
BUSIERE	80	NP B174 1	A. Busiere <i>et al.</i>	(BGNA, SACL, LAPP)
MARINELLI	80B	PL 94B 433	M. Marinelli, G. Morpurgo	(GENO)
		Also	M. Marinelli, G. Morpurgo	(GENO)
BOYD	79	PRL 43 1288	R.N. Boyd <i>et al.</i>	(OSU)
BOZZOLI	79	NP B159 363	W. Bozzoli <i>et al.</i>	(BGNA, LAPP, SACL+)

See key on page 601

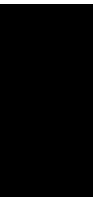
# Quark Particle Listings

## Free Quark Searches

LARUE	79	PRL 42 142	G.S. Larue, W.M. Fairbank, J.D. Phillips	(STAN)	MORPURGO	70	NIM 79 95	G. Morpurgo, G. Gallinaro, G. Palmieri	(GENO)
Also		PRL 42 1019	G.S. Larue, W.M. Fairbank, J.D. Phillips		ALLABY	69B	NC 64A 75	J.V. Allaby <i>et al.</i>	(CERN)
OGOROD...	79	JETP 49 953	D.D. Ogorodnikov, I.M. Samoilov, A.M. Solntsev		ANTIPOV	69	PL 29B 245	Y.M. Antipov <i>et al.</i>	(SERP)
		Translated from ZETF 76 1881.			ANTIPOV	69B	PL 30B 576	Y.M. Antipov <i>et al.</i>	(SERP)
STEVENSON	79	PR D20 82	M.L. Stevenson	(LBL)	CAIRNS	69	PR 186 1394	I. Cairns <i>et al.</i>	(SYDN)
BASILE	78	NC 45A 171	M. Basile <i>et al.</i>	(CERN, BGNA)	COOK	69	PR 188 2092	D.D. Cook <i>et al.</i>	(ILL)
BASILE	78B	NC 45A 201	M. Basile <i>et al.</i>	(CERN, BGNA)	FUKUSHIMA	69	PR 178 2058	Y. Fukushima <i>et al.</i>	(TOKY)
BOYD	78	PRL 40 216	R.N. Boyd <i>et al.</i>	(ROCH)	MCCUSKER	69	PRL 23 658	C.B.A. McCusker, I. Cairns	(SYDN)
BOYD	78B	PL 72B 484	R.N. Boyd <i>et al.</i>	(ROCH)	BELLAMY	68	PR 166 1391	E.H. Bellamy <i>et al.</i>	(STAN, SLAC)
LUND	78	RA 25 75	T. Lund, R. Brandt, Y. Fares	(MARB)	BJORNBOE	68	NC B53 241	J. Bjornboe <i>et al.</i>	(BOHR, TATA, BERN+)
PUTT	78	PR D17 1466	G.D. Putt, P.C.M. Yock	(AUCK)	BRAGINSK	68	JETP 27 51	V.B. Braginsky <i>et al.</i>	(MOSU)
SCHIFFER	78	PR D17 2241	J.P. Schiffer <i>et al.</i>	(CHIC, ANL)			Translated from ZETF 54 91.		
YOCK	78	PR D18 641	P.C.M. Yock	(AUCK)	BRIATORE	68	NC 57A 850	L. Briatore <i>et al.</i>	(TORI, CERN, BGNA)
ANTREASYAN	77	PRL 39 513	D. Antreasyan <i>et al.</i>	(EFI, PRIN)	FRANZINI	68	PRL 21 1013	P. Franzini, S. Shulman	(COLU)
BASILE	77	NC 40A 41	M. Basile <i>et al.</i>	(CERN, BGNA)	GARMIRE	68	PR 166 1280	G. Garmire, C. Leong, V. Sreekantan	(MIT)
BLAND	77	PRL 39 369	R.W. Bland <i>et al.</i>	(SFSU)	HANAYAMA	68	CJP 46 5734	Y. Hanayama <i>et al.</i>	(OSAK)
GALLINARO	77	PRL 38 1235	G. Gallinaro, M. Marinelli, G. Morpurgo	(GENO)	KASHA	68	PR 172 1297	H. Kasha, R.J. Stefanski	(BNL, YALE)
JONES	77B	RMP 49 717	L.W. Jones		KASHA	68B	PRL 20 217	H. Kasha <i>et al.</i>	(BNL, YALE)
LARUE	77	PRL 38 1011	G.S. Larue, W.M. Fairbank, A.F. Hebard	(STAN)	KASHA	68C	CJP 46 5730	H. Kasha <i>et al.</i>	(BNL, YALE)
MULLER	77	SCI 196 521	R.A. Muller <i>et al.</i>	(LBL)	RANK	68	PR 176 1635	D. Rank	(MICH)
OGOROD...	77	JETP 45 857	D.D. Ogorodnikov, I.M. Samoilov, A.M. Solntsev		BARTON	67	PRSL 90 87	J.C. Barton	(NPOL)
		Translated from ZETF 72 1633.			BATHOW	67	PL 25B 163	G. Bathow <i>et al.</i>	(DESY)
BALDIN	76	SJNP 22 264	B.Y. Baldin <i>et al.</i>	(JINR)	BUHLER	67	NC 49A 209	A. Buhler-Broglin <i>et al.</i>	(CERN, BGNA)
		Translated from YAF 22 512.			BUHLER	67B	NC 51A 837	A. Buhler-Broglin <i>et al.</i>	(CERN, BGNA+)
BRIATORE	76	NC 31A 553	L. Briatore <i>et al.</i>	(LCGT, FRAS, FREIB)	FOSS	67	PL 25B 166	J. Foss <i>et al.</i>	(MIT)
STEVENS	76	PR D14 716	C.M. Stevens, J.P. Schiffer, W. Chupka	(ANL)	GOMEZ	67	PRL 18 1022	R. Gomez <i>et al.</i>	(CIT)
ALBROW	75	NP B97 189	M.G. Albrow <i>et al.</i>	(CERN, DARE, FOM+)	KASHA	67	PR 154 1263	H. Kasha <i>et al.</i>	(BNL, YALE)
FABJAN	75	NP B101 349	C.W. Fabjan <i>et al.</i>	(CERN, MPIM)	STOVER	67	PR 164 1599	R.W. Stover, T.I. Moran, J.W. Trischka	(SYRA)
HAZEN	75	NP B95 189	W.E. Hazen <i>et al.</i>	(MICH, LEED)	BARTON	66	PL 21 360	J.C. Barton, C.T. Stockel	(NPOL)
JOVANOV...	75	PL 56B 105	J.V. Jovanovich <i>et al.</i>	(MANI, AACH, CERN+)	BENNETT	66	PRL 17 1196	W.R. Bennett	(YALE)
KRISOR	75	NC 27A 132	K. Krisor	(AACH3)	BUHLER	66	NC 45A 620	A. Buhler-Broglin <i>et al.</i>	(CERN, BGNA+)
CLARK	74B	PR D10 2721	A.F. Clark <i>et al.</i>	(LLL)	CHUPKA	66	PRL 17 60	W.A. Chupka, J.P. Schiffer, C.M. Stevens	(ANL)
GALIK	74	PR D9 1856	R.S. Galik <i>et al.</i>	(SLAC, FNAL)	GALLINARO	66	PL 23 609	G. Gallinaro, G. Morpurgo	(GENO)
KIFUNE	74	JPS J 36 629	T. Kifune <i>et al.</i>	(TOKY, KEK)	KASHA	66	PR 150 1140	H. Kasha, L.B. Leipuner, R.K. Adair	(BNL, YALE)
NASH	74	PRL 32 858	T. Nash <i>et al.</i>	(FNAL, CORN, NYU)	LAMB	66	PRL 17 1068	R.C. Lamb <i>et al.</i>	(ANL)
ALPER	73	PL 46B 265	B. Alper <i>et al.</i>	(CERN, LVP, LUND, BOHR+)	DELISE	65	PR 140B 458	D.A. de Lise, T. Bowen	(ARIZ)
ASHTON	73	JP A6 577	F. Ashton <i>et al.</i>	(DURH)	DORFAN	65	PRL 14 999	D.E. Dorfan <i>et al.</i>	(COLU)
HICKS	73B	NC 14A 65	R.B. Hicks, R.W. Flint, S. Standil	(MANI)	FRANZINI	65B	PRL 14 196	P. Franzini <i>et al.</i>	(BNL, COLU)
LEIPUNER	73	PRL 31 1226	L.B. Leipuner <i>et al.</i>	(BNL, YALE)	MASSAM	65	NC 40A 589	T. Massam, T. Muller, A. Z'ichichi	(CERN)
BEAUCHAMP	72	PR D6 1211	W.T. Beauchamp <i>et al.</i>	(ARIZ)	BINGHAM	64	PL 9 201	H.H. Bingham <i>et al.</i>	(CERN, EPOL)
BOHM	72B	PRL 28 326	A. Bohm <i>et al.</i>	(AACH)	BLUM	64	PRL 13 353A	W. Blum <i>et al.</i>	(CERN)
BOTT	72	PL 40B 693	M. Bott-Bodenhausen <i>et al.</i>	(CERN, MPIM)	BOWEN	64	PRL 13 728	T. Bowen <i>et al.</i>	(ARIZ)
COX	72	PR D6 1203	A.J. Cox <i>et al.</i>	(ARIZ)	HAGOPIAN	64	PRL 13 280	V. Hagopian <i>et al.</i>	(PENN, BNL)
CROUCH	72	PR D5 2667	M.F. Crouch, K. Mori, G.R. Smith	(CASE)	LEIPUNER	64	PRL 12 423	L.B. Leipuner <i>et al.</i>	(BNL, YALE)
DARDO	72	NC 9A 319	M. Dardo <i>et al.</i>	(TORI)	MORRISON	64	PL 9 199	D.R.O. Morrison	(CERN)
EVANS	72	PRSE A70 143	G.R. Evans <i>et al.</i>	(EDIN, LEED)	SUNYAR	64	PR 136 B1157	A.W. Sunyar, A.Z. Schwarzschild, P.I. Connors	(BNL)
TONWAR	72	JP A5 569	S.C. Tonwar, S. Naranan, B.V. Sreekantan	(TATA)	HILLAS	59	NAT 184 B92	A.M. Hillas, T.E. Cranshaw	(AERE)
ANTIPOV	71	NP B29 374	Y.M. Antipov <i>et al.</i>	(SERP)	MILLIKAN	10	Phil Mag 19 209	R.A. Millikan	(CHIC)
CHIN	71	NC 2A 419	S. Chin <i>et al.</i>	(OSAK)					
CLARK	71B	PRL 27 51	A.F. Clark <i>et al.</i>	(LLL, LBL)					
HAZEN	71	PRL 26 582	W.E. Hazen	(MICH)					
BOSIA	70	NC 66A 167	G.F. Bosia, L. Briatore	(TORI)					
CHU	70	PRL 24 917	W.T. Chu <i>et al.</i>	(OSU, ROSE, KANS)					
Also		PRL 25 550	W.W.M. Allison <i>et al.</i>	(ANL)					
ELBERT	70	NP B20 217	J.W. Elbert <i>et al.</i>	(WISC)					
FAISSNER	70B	PRL 24 1357	H. Faissner <i>et al.</i>	(AACH3)					
KRIDER	70	PR D1 835	E.P. Krider, T. Bowen, R.M. Kalbach	(ARIZ)					

### OTHER RELATED PAPERS

LYONS	85	PRPL C129 225	L. Lyons	(OXF)
Review				
MARINELLI	82	PRPL 85 161	M. Marinelli, G. Morpurgo	(GENO)
Review				



# **LIGHT UNFLAVORED MESONS ( $S = C = B = 0$ )**

• $\pi^\pm$	849
• $\pi^0$	853
• $\eta$	855
• $f_0(500)$	861
• $\rho(770)$	870
• $\omega(782)$	876
• $\eta'(958)$	881
• $f_0(980)$	886
• $a_0(980)$	889
• $\phi(1020)$	891
• $h_1(1170)$	897
• $b_1(1235)$	898
• $a_1(1260)$	899
• $f_2(1270)$	901
• $f_1(1285)$	904
• $\eta(1295)$	907
• $\pi(1300)$	907
• $a_2(1320)$	908
• $f_0(1370)$	911
• $h_1(1380)$	914
• $\pi_1(1400)$	914
• $\eta(1405)$	915
• $f_1(1420)$	920
• $\omega(1420)$	921
• $f_2(1430)$	922
• $a_0(1450)$	923
• $\rho(1450)$	924
• $\eta(1475)$	926
• $f_0(1500)$	927
• $f_1(1510)$	930
• $f_2'(1525)$	931
• $f_2(1565)$	933
• $\rho(1570)$	934
• $h_1(1595)$	935
• $\pi_1(1600)$	935
• $a_1(1640)$	936
• $f_2(1640)$	936
• $\eta_2(1645)$	937
• $\omega(1650)$	937
• $\omega_3(1670)$	938
• $\pi_2(1670)$	939
• $\phi(1680)$	940
• $\rho_3(1690)$	942
• $\rho(1700)$	945
• $a_2(1700)$	949
• $f_0(1710)$	950
• $\eta(1760)$	953
• $\pi(1800)$	953
• $f_2(1810)$	954
• $X(1835)$	955
• $X(1840)$	956
• $a_1(1420)$	957
• $\phi_3(1850)$	957
• $\eta_2(1870)$	957
• $\pi_2(1880)$	958
• $\rho(1900)$	958
• $f_2(1910)$	958
• $a_0(1950)$	959
• $f_2(1950)$	960
• $\rho_3(1990)$	961
• $f_2(2010)$	961
• $f_0(2020)$	961

• $a_4(2040)$	962
• $f_4(2050)$	962
• $\pi_2(2100)$	964
• $f_0(2100)$	964
• $f_2(2150)$	965
• $\rho(2150)$	966
• $\phi(2170)$	967
• $f_0(2200)$	968
• $f_J(2220)$	968
• $\eta(2225)$	969
• $\rho_3(2250)$	970
• $f_2(2300)$	970
• $f_4(2300)$	971
• $f_0(2330)$	971
• $f_2(2340)$	972
• $\rho_5(2350)$	972
• $a_6(2450)$	973
• $f_6(2510)$	973

## **OTHER LIGHT UNFLAVORED ( $S = C = B = 0$ )**

Further States	974
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## **STRANGE MESONS ( $S = \pm 1, C = B = 0$ )**

• $K^\pm$	979
• $K^0$	998
• $K_S^0$	1002
• $K_L^0$	1006
• $K_0^*(800)$	1028
• $K^*(892)$	1029
• $K_1(1270)$	1031
• $K_1(1400)$	1033
• $K^*(1410)$	1033
• $K_0^*(1430)$	1034
• $K_2^*(1430)$	1035
• $K(1460)$	1037
• $K_2(1580)$	1037
• $K(1630)$	1037
• $K_1(1650)$	1038
• $K^*(1680)$	1038
• $K_2(1770)$	1038
• $K_3^*(1780)$	1039
• $K_2(1820)$	1040
• $K(1830)$	1041
• $K_0^*(1950)$	1041
• $K_2^*(1980)$	1041
• $K_4^*(2045)$	1041
• $K_2(2250)$	1042
• $K_3(2320)$	1042
• $K_5^*(2380)$	1043
• $K_4(2500)$	1043
• $K(3100)$	1043

## **CHARMED MESONS ( $C = \pm 1$ )**

• $D^\pm$	1044
• $D^0$	1061
• $D^*(2007)^0$	1095
• $D^*(2010)^\pm$	1095
• $D_0^*(2400)^0$	1096
• $D_0^*(2400)^\pm$	1097
• $D_1(2420)^0$	1097
• $D_1(2420)^\pm$	1098
• $D_1(2430)^0$	1099
• $D_2^*(2460)^0$	1099
• $D_2^*(2460)^\pm$	1101

• Indicates the particle is in the Meson Summary Table

(continued on the next page)

$D(2550)^0$	1101	$X(3940)$	1444
$D_J^*(2600)$	1102	$X(4020)$	1445
$D^*(2640)^\pm$	1102	• $\psi(4040)$	1445
$D(2740)^0$	1102	$X(4050)^\pm$	1447
$D(2750)$	1103	$X(4055)^\pm$	1448
$D(3000)^0$	1103	$X(4140)$	1448
<b>CHARMED, STRANGE MESONS (<math>C = S = \pm 1</math>)</b>		• $\psi(4160)$	1449
• $D_s^\pm$	1104	$X(4160)$	1451
• $D_s^{*\pm}$	1130	$X(4200)^\pm$	1451
• $D_{s0}^*(2317)^\pm$	1131	$X(4230)$	1451
• $D_{s1}(2460)^\pm$	1131	$X(4240)^\pm$	1452
• $D_{s1}(2536)^\pm$	1133	$X(4250)^\pm$	1452
• $D_{s2}(2573)$	1134	• $X(4260)$	1452
• $D_{s1}^*(2700)^\pm$	1135	$X(4350)$	1455
$D_{s1}^*(2860)^\pm$	1135	$X(4360)$	1455
$D_{s3}^*(2860)^\pm$	1136	• $\psi(4415)$	1456
$D_{sJ}(3040)^\pm$	1136	$X(4430)^\pm$	1457
		$X(4660)$	1458
<b>BOTTOM MESONS (<math>B = \pm 1</math>)</b>		<b><math>b\bar{b}</math> MESONS</b>	
$B$ -particle organization	1137	Bottomonium system	1460
• $B^\pm$	1150	$\eta_b(1S)$	1461
• $B^0$	1206	• $\Upsilon(1S)$	1462
• $B^\pm/B^0$ ADMIXTURE	1285	• $\chi_{b0}(1P)$	1467
• $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE	1305	• $\chi_{b1}(1P)$	1468
$V_{cb}$ and $V_{ub}$ CKM Matrix Elements	1313	• $h_b(1P)$	1470
• $B^*$	1328	• $\chi_{b2}(1P)$	1470
• $B_1(5721)^+$	1328	$\eta_b(2S)$	1472
• $B_1(5721)^0$	1329	• $\Upsilon(2S)$	1472
$B_J^*(5732)$	1329	$\Upsilon(1D)$	1476
• $B_2^*(5747)^+$	1329	• $\chi_{b0}(2P)$	1477
• $B_2^*(5747)^0$	1330	• $\chi_{b1}(2P)$	1478
$B_J(5840)^+$	1330	$h_b(2P)$	1480
$B_J(5840)^0$	1331	• $\chi_{b2}(2P)$	1480
$B_J(5970)^+$	1331	• $\Upsilon(3S)$	1482
$B_J(5970)^0$	1332	$\chi_{b1}(3P)$	1486
<b>BOTTOM, STRANGE MESONS (<math>B = \pm 1, S = \mp 1</math>)</b>		• $\Upsilon(4S)$	1486
• $B_s^0$	1333	$X(10610)^\pm$	1488
• $B_s^*$	1351	$X(10610)^0$	1489
• $B_{s1}(5830)^0$	1351	$X(10650)^\pm$	1489
• $B_{s2}^*(5840)^0$	1351	• $\Upsilon(10860)$	1490
$B_{sJ}^*(5850)$	1352	• $\Upsilon(11020)$	1492
<b>BOTTOM, CHARMED MESONS (<math>B = C = \pm 1</math>)</b>		<b>Notes in the Meson Listings</b>	
• $B_c^+$	1353	Form factors for radiative pion & kaon decays (rev.)	850
$B_c(2S)^\pm$	1355	Note on scalar mesons below 2 GeV (rev.)	861
<b><math>c\bar{c}</math> MESONS</b>		Pseudo-scalar & -vector mesons around 1400 MeV (rev.)	915
Charmonium system	1364	The $\rho(1450)$ and the $\rho(1700)$ (rev.)	945
• $\eta_c(1S)$	1364	Rare kaon decays (rev.)	981
• $J/\psi(1S)$	1371	Dalitz plot parameters for $K \rightarrow 3\pi$ decays	992
• $\chi_{c0}(1P)$	1390	$K_{\ell 3}^\pm$ and $K_{\ell 3}^0$ form factors	993
• $\chi_{c1}(1P)$	1400	$CPT$ Invariance tests in neutral kaon decay	999
• $h_c(1P)$	1407	$V_{ud}$ , $V_{us}$ , Cabibbo angle, and CKM unitarity (rev.)	1011
• $\chi_{c2}(1P)$	1408	$CP$ -Violation in $K_L$ decays	1019
• $\eta_c(2S)$	1417	Review of multibody charm analyses (rev.)	1048
• $\psi(2S)$	1419	$D^0$ - $\bar{D}^0$ mixing (rev.)	1061
• $\psi(3770)$	1434	$D_s^+$ branching fractions (rev.)	1106
$\psi(3823)$	1439	Leptonic decays of charged pseudoscalar mesons (rev.)	1109
• $X(3872)$	1440	Production and decay of $b$ -flavored hadrons (rev.)	1137
• $X(3900)$	1442	A note on HFAG activities (rev.)	1149
• $X(3915)$	1443	Polarization in $B$ decays (rev.)	1252
$\chi_{c2}(2P)$	1444	$B^0$ - $\bar{B}^0$ mixing (rev.)	1259
• Indicates the particle is in the Meson Summary Table		Semileptonic $B$ decays, $V_{cb}$ and $V_{ub}$ (rev.)	1313
		Heavy quarkonium spectroscopy (rev.)	1355
		Non- $q\bar{q}$ candidates (rev.)	1494



# LIGHT UNFLAVORED MESONS ( $S = C = B = 0$ )

For  $I = 1$  ( $\pi, b, \rho, a$ ):  $u\bar{d}, (u\bar{u}-d\bar{d})/\sqrt{2}, d\bar{u}$ ;  
for  $I = 0$  ( $\eta, \eta', h, h', \omega, \phi, f, f'$ ):  $c_1(u\bar{u}+d\bar{d}) + c_2(s\bar{s})$

 $\pi^\pm$ 

$$I^G(J^P) = 1^-(0^-)$$

We have omitted some results that have been superseded by later experiments. The omitted results may be found in our 1988 edition Physics Letters **B204** 1 (1988).

## $\pi^\pm$ MASS

The most accurate charged pion mass measurements are based upon x-ray wavelength measurements for transitions in  $\pi^-$ -mesonic atoms. The observed line is the blend of three components, corresponding to different K-shell occupancies. JECKELMANN 94 revisits the occupancy question, with the conclusion that two sets of occupancy ratios, resulting in two different pion masses (Solutions A and B), are equally probable. We choose the higher Solution B since only this solution is consistent with a positive mass-squared for the muon neutrino, given the precise muon momentum measurements now available (DAUM 91, ASSAMAGAN 94, and ASSAMAGAN 96) for the decay of pions at rest. Earlier mass determinations with pi-mesonic atoms may have used incorrect K-shell screening corrections.

Measurements with an error of  $> 0.005$  MeV have been omitted from this Listing.

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>139.57018±0.00035 OUR FIT</b>	Error includes scale factor of 1.2.				
<b>139.57018±0.00035 OUR AVERAGE</b>	Error includes scale factor of 1.2.				
139.57071±0.00053	<sup>1</sup> LENZ	98	CNTR	— pionic N2-atoms gas target	
139.56995±0.00035	<sup>2</sup> JECKELMANN 94	CNTR	—	$\pi^-$ atom, Soln. B	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
139.57022±0.00014	<sup>3</sup> ASSAMAGAN 96	SPEC	+	$\pi^+ \rightarrow \mu^+ \nu_\mu$	
139.56782±0.00037	<sup>4</sup> JECKELMANN 94	CNTR	—	$\pi^-$ atom, Soln. A	
139.56996±0.00067	<sup>5</sup> DAUM	91	SPEC	+	$\pi^+ \rightarrow \mu^+ \nu$
139.56752±0.00037	<sup>6</sup> JECKELMANN 86b	CNTR	—	Mesonic atoms	
139.5704 ±0.0011	<sup>5</sup> ABELA	84	SPEC	+	See DAUM 91
139.5664 ±0.0009	<sup>7</sup> LU	80	CNTR	—	Mesonic atoms
139.5686 ±0.0020	CARTER	76	CNTR	—	Mesonic atoms
139.5660 ±0.0024	<sup>7,8</sup> MARUSHEN... 76	CNTR	—	Mesonic atoms	

<sup>1</sup> LENZ 98 result does not suffer K-electron configuration uncertainties as does JECKELMANN 94.

<sup>2</sup> JECKELMANN 94 Solution B (dominant 2-electron K-shell occupancy), chosen for consistency with positive  $m_{\nu_\mu}^2$ .

<sup>3</sup> ASSAMAGAN 96 measures the  $\mu^+$  momentum  $p_\mu$  in  $\pi^+ \rightarrow \mu^+ \nu_\mu$  decay at rest to be  $29.79200 \pm 0.00011$  MeV/c. Combined with the  $\mu^+$  mass and the assumption  $m_{\nu_\mu} = 0$ , this gives the  $\pi^+$  mass above; if  $m_{\nu_\mu} > 0$ ,  $m_{\pi^+}$  given above is a lower limit.

Combined instead with  $m_\mu$  and (assuming  $CPT$ ) the  $\pi^-$  mass of JECKELMANN 94,  $p_\mu$  gives an upper limit on  $m_{\nu_\mu}$  (see the  $\nu_\mu$ ).

<sup>4</sup> JECKELMANN 94 Solution A (small 2-electron K-shell occupancy) in combination with either the DAUM 91 or ASSAMAGAN 94 pion decay muon momentum measurement yields a significantly negative  $m_{\nu_\mu}^2$ . It is accordingly not used in our fits.

<sup>5</sup> The DAUM 91 value includes the ABELA 84 result. The value is based on a measurement of the  $\mu^+$  momentum for  $\pi^+$  decay at rest,  $p_\mu = 29.79179 \pm 0.00053$  MeV, uses  $m_\mu = 105.658389 \pm 0.000034$  MeV, and assumes that  $m_{\nu_\mu} = 0$ . The last assumption means that in fact the value is a lower limit.

<sup>6</sup> JECKELMANN 86b gives  $m_{\pi^+}/m_e = 273.12677(71)$ . We use  $m_e = 0.51099906(15)$  MeV from COHEN 87. The authors note that two solutions for the probability distribution of K-shell occupancy fit equally well, and use other data to choose the lower of the two possible  $\pi^\pm$  masses.

<sup>7</sup> These values are scaled with a new wavelength-energy conversion factor  $V\lambda = 1.23984244(37) \times 10^{-6}$  eV m from COHEN 87. The LU 80 screening correction relies upon a theoretical calculation of inner-shell refilling rates.

<sup>8</sup> This MARUSHENKO 76 value used at the authors' request to use the accepted set of calibration  $\gamma$  energies. Error increased from 0.0017 MeV to include QED calculation error of 0.0017 MeV (12 ppm).

$$m_{\pi^+} - m_{\mu^+}$$

Measurements with an error  $> 0.05$  MeV have been omitted from this Listing.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
33.91157±0.00067	1	DAUM	91	SPEC	+	$\pi^+ \rightarrow \mu^+ \nu$
33.9111 ± 0.0011		ABELA	84	SPEC		See DAUM 91
33.925 ± 0.025		BOOTH	70	CNTR	+	Magnetic spect.
33.881 ± 0.035	145	HYMAN	67	HEBC	+	$K^-$ He

<sup>1</sup> The DAUM 91 value assumes that  $m_{\nu_\mu} = 0$  and uses our  $m_\mu = 105.658389 \pm 0.000034$  MeV.

$$(m_{\pi^+} - m_{\pi^-}) / m_{\text{average}}$$

A test of  $CPT$  invariance.

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN
<b>2 ± 5</b>	AYRES	71 CNTR

## $\pi^\pm$ MEAN LIFE

Measurements with an error  $> 0.02 \times 10^{-8}$  s have been omitted.

VALUE ( $10^{-8}$ s)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>2.6033 ± 0.0005 OUR AVERAGE</b>	Error includes scale factor of 1.2.				
2.60361 ± 0.00052	<sup>1</sup> KOPTEV	95	SPEC	+	Surface $\mu^+$ 's
2.60231 ± 0.00050 ± 0.00084	NUMAO	95	SPEC	+	Surface $\mu^+$ 's
2.609 ± 0.008	DUNAITSEV	73	CNTR	+	
2.602 ± 0.004	AYRES	71	CNTR	±	
2.604 ± 0.005	NORDBERG	67	CNTR	+	
2.602 ± 0.004	ECKHAUSE	65	CNTR	+	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.640 ± 0.008	<sup>2</sup> KINSEY	66	CNTR	+	

<sup>1</sup> KOPTEV 95 combines the statistical and systematic errors; the statistical error dominates.

<sup>2</sup> Systematic errors in the calibration of this experiment are discussed by NORDBERG 67.

$$(\tau_{\pi^+} - \tau_{\pi^-}) / \tau_{\text{average}}$$

A test of  $CPT$  invariance.

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN
<b>5.5 ± 7.1</b>	AYRES	71 CNTR
• • • We do not use the following data for averages, fits, limits, etc. • • •		
−14 ± 29	PETRUKHIN	68 CNTR
40 ± 70	BARDON	66 CNTR
23 ± 40	<sup>1</sup> LOBKOWICZ	66 CNTR

<sup>1</sup> This is the most conservative value given by LOBKOWICZ 66.

## $\pi$ ELECTRIC POLARIZABILITY $\alpha_\pi$

See HOLSTEIN 14 for a general review on hadron polarizability.

VALUE ( $10^{-4}$ fm <sup>3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.0 ± 0.6 ± 0.7</b>	63k	<sup>1</sup> ADOLPH	15A	SPEC $\pi^- \gamma \rightarrow \pi^- \gamma$ Compton scatt.

<sup>1</sup> Value is derived assuming  $\alpha_\pi = -\beta_\pi$ .

## $\pi^\pm$ DECAY MODES

$\pi^-$  modes are charge conjugates of the modes below.

For decay limits to particles which are not established, see the section on Searches for Axions and Other Very Light Bosons.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $\mu^+ \nu_\mu$	[a] (99.98770 ± 0.00004) %	
$\Gamma_2$ $\mu^+ \nu_\mu \gamma$	[b] ( 2.00 ± 0.25 ) × $10^{-4}$	
$\Gamma_3$ $e^+ \nu_e$	[a] ( 1.230 ± 0.004 ) × $10^{-4}$	
$\Gamma_4$ $e^+ \nu_e \gamma$	[b] ( 7.39 ± 0.05 ) × $10^{-7}$	
$\Gamma_5$ $e^+ \nu_e \pi^0$	( 1.036 ± 0.006 ) × $10^{-8}$	
$\Gamma_6$ $e^+ \nu_e e^+ e^-$	( 3.2 ± 0.5 ) × $10^{-9}$	
$\Gamma_7$ $e^+ \nu_e \nu \bar{\nu}$	< 5	× $10^{-6}$ 90%

## Lepton Family number (LF) or Lepton number (L) violating modes

$\Gamma_8$ $\mu^+ \bar{\nu}_e$	L	[c] < 1.5	× $10^{-3}$ 90%
$\Gamma_9$ $\mu^+ \nu_e$	LF	[c] < 8.0	× $10^{-3}$ 90%
$\Gamma_{10}$ $\mu^- e^+ e^+ \nu$	LF	< 1.6	× $10^{-6}$ 90%

[a] Measurements of  $\Gamma(e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$  always include decays with  $\gamma$ 's, and measurements of  $\Gamma(e^+ \nu_e \gamma)$  and  $\Gamma(\mu^+ \nu_\mu \gamma)$  never include low-energy  $\gamma$ 's. Therefore, since no clean separation is possible, we consider the modes with  $\gamma$ 's to be subreactions of the modes without them, and let  $[\Gamma(e^+ \nu_e) + \Gamma(\mu^+ \nu_\mu)]/\Gamma_{\text{total}} = 100\%$ .

[b] See the Particle Listings below for the energy limits used in this measurement; low-energy  $\gamma$ 's are not included.

[c] Derived from an analysis of neutrino-oscillation experiments.

# Meson Particle Listings

 $\pi^\pm$ 

$\pi^+$ BRANCHING RATIOS	
$\Gamma(e^+ \nu_e)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$

See note [a] in the list of  $\pi^+$  decay modes just above, and see also the next block of data. See also the note on “Decay Constants of Charged Pseudoscalar Mesons” in the  $D_s^+$  Listings.

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b><math>1.230 \pm 0.004</math> OUR EVALUATION</b>	

$[\Gamma(e^+ \nu_e) + \Gamma(e^+ \nu_e \gamma)]/[\Gamma(\mu^+ \nu_\mu) + \Gamma(\mu^+ \nu_\mu \gamma)]$	$(\Gamma_3 + \Gamma_4)/(\Gamma_1 + \Gamma_2)$
---	---

See note [a] in the list of  $\pi^+$  decay modes above. See NUMAO 92 for a discussion of  $e$ - $\mu$  universality. See also the note on “Decay Constants of Charged Pseudoscalar Mesons” in the  $D_s^+$  Listings.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1.2327 \pm 0.0023</math> OUR AVERAGE</b>					
$1.2344 \pm 0.0023 \pm 0.0019$	400k	AGUILAR-AR...15	CNTR	+	Stopping $\pi^+$
$1.2346 \pm 0.0035 \pm 0.0036$	120k	CZAPEK 93	CALO		Stopping $\pi^+$
$1.2265 \pm 0.0034 \pm 0.0044$	190k	BRITTON 92	CNTR		Stopping $\pi^+$
$1.218 \pm 0.014$	32k	BRYMAN 86	CNTR		Stopping $\pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.273 \pm 0.028$	11k	<sup>1</sup> DICAPUA 64	CNTR		
$1.21 \pm 0.07$		ANDERSON 60	SPEC		

<sup>1</sup> DICAPUA 64 has been updated using the current mean life.

$\Gamma(\mu^+ \nu_\mu)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
Note that measurements here do not cover the full kinematic range.	

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>2.0 \pm 0.24 \pm 0.08</math></b>		<sup>1</sup> BRESSI 98	CALO	+	Stopping $\pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.24 \pm 0.25$	26	CASTAGNOLI 58	EMUL		$KE_\mu < 3.38$ MeV

<sup>1</sup> BRESSI 98 result is given for  $E_\gamma > 1$  MeV only. Result agrees with QED expectation,  $2.283 \times 10^{-4}$  and does not confirm discrepancy of earlier experiment CASTAGNOLI 58.

$\Gamma(e^+ \nu_e \gamma)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
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The very different values reflect the very different kinematic ranges covered (bigger range, bigger value). And none of them covers the whole kinematic range.

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>73.86 \pm 0.54</math></b>	65k	<sup>1</sup> BYCHKOV 09	PIBE		$e^+ \nu \gamma$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$16.1 \pm 2.3$		<sup>2</sup> BOLOTOV 90B	SPEC		17 GeV $\pi^- \rightarrow e^- \bar{\nu}_e \gamma$
$5.6 \pm 0.7$	226	<sup>3</sup> STETZ 78	SPEC		$P_e > 56$ MeV/c
3.0	143	DEPOMMIER 63B	CNTR		(KE) $e^+ \gamma > 48$ MeV

<sup>1</sup> This BYCHKOV 09 value is for  $E_\gamma > 10$  MeV and  $\Theta_{e^+ \gamma} > 40^\circ$ .

<sup>2</sup> BOLOTOV 90B is for  $E_\gamma > 21$  MeV,  $E_e > 70 - 0.8 E_\gamma$ .

<sup>3</sup> STETZ 78 is for an  $e^- \gamma$  opening angle  $> 132^\circ$ . Obtains 3.7 when using same cutoffs as DEPOMMIER 63B.

$\Gamma(e^+ \nu_e \pi^0)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
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VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1.036 \pm 0.006</math> OUR AVERAGE</b>					
$1.036 \pm 0.006$	64k	<sup>1,2</sup> POCANIC 04	PIBE	+	$\pi$ decay at rest
$1.026 \pm 0.039$	1224	<sup>3</sup> MCFARLANE 85	CNTR	+	Decay in flight
$1.00 \pm 0.08$	332	DEPOMMIER 68	CNTR	+	
$1.07 \pm 0.21$	38	<sup>4</sup> BACASTOW 65	OSPK	+	
$1.10 \pm 0.26$		<sup>4</sup> BERTRAM 65	OSPK	+	
$1.1 \pm 0.2$	43	<sup>4</sup> DUNAITSEV 65	CNTR	+	
$0.97 \pm 0.20$	36	<sup>4</sup> BARTLETT 64	OSPK	+	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.15 \pm 0.22$  52 <sup>4</sup>DEPOMMIER 63 CNTR + See DEPOMMIER 68

<sup>1</sup> POCANIC 04 normalizes to  $e^+ \nu_e$  decays, using the PDG 2004 value  $B(\pi^+ \rightarrow e^+ \nu_e) = (1.230 \pm 0.004) \times 10^{-4}$ . We add their statistical  $(0.004 \times 10^{-8})$ , systematic  $(0.004 \times 10^{-8})$  and systematic error due to the uncertainty of  $B(\pi^+ \rightarrow e^+ \nu_e)$   $(0.003 \times 10^{-8})$  in quadrature.

<sup>2</sup> This result can be used to calculate  $V_{ud}$  from pion beta decay:  $V_{ud}^{PIBETA} = 0.9728 \pm 0.0030$ .

<sup>3</sup> MCFARLANE 85 combines a measured rate  $(0.394 \pm 0.015)/s$  with 1982 PDG mean life.

<sup>4</sup> DEPOMMIER 68 says the result of DEPOMMIER 63 is at least 10% too large because of a systematic error in the  $\pi^0$  detection efficiency, and that this may be true of all the previous measurements (also V. Soergel, private communication, 1972).

$\Gamma(e^+ \nu_e e^+ e^-)/\Gamma(\mu^+ \nu_\mu)$	$\Gamma_6/\Gamma_1$
---	---------------------

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.2 \pm 0.5 \pm 0.2</math></b>		98	EGLI 89	SPEC	Uses $R_{PCAC} = 0.068 \pm 0.004$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.46 \pm 0.16 \pm 0.07$		7	<sup>1</sup> BARANOV 92	SPEC	Stopped $\pi^+$
$< 4.8$	90		KORENCHE... 76B	SPEC	
$< 34$	90		KORENCHE... 71	OSPK	

<sup>1</sup> This measurement by BARANOV 92 is of the structure-dependent part of the decay. The value depends on values assumed for ratios of form factors.

$\Gamma(e^+ \nu_e \nu \bar{\nu})/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
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VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN
<b><math>&lt; 5</math></b>	90	PICCIOTTO 88	SPEC

$\Gamma(\mu^+ \bar{\nu}_e)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
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Forbidden by total lepton number conservation. See the note on “Decay Constants of Charged Pseudoscalar Mesons” in the  $D_s^+$  Listings.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.5</math></b>	90	<sup>1</sup> COOPER 82	HLBC	Wideband $\nu$ beam

<sup>1</sup> COOPER 82 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton number violation.

$\Gamma(\mu^+ \nu_e)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
---	-------------------

Forbidden by lepton family number conservation.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 8.0</math></b>	90	<sup>1</sup> COOPER 82	HLBC	Wideband $\nu$ beam

<sup>1</sup> COOPER 82 limit on  $\nu_e$  observation is here interpreted as a limit on lepton family number violation.

$\Gamma(\mu^- e^+ e^+ \nu)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
---	----------------------

Forbidden by lepton family number conservation.

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	CHG
<b><math>&lt; 1.6</math></b>	90	BARANOV 91B	SPEC	+
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 7.7$	90	KORENCHE... 87	SPEC	+

## $\pi^+$ — POLARIZATION OF EMITTED $\mu^+$

 $\pi^+ \rightarrow \mu^+ \nu$ 

Tests the Lorentz structure of leptonic charged weak interactions.

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< (-0.9959)$	90	<sup>1</sup> FETSCHER 84	RVUE	+	
$-0.99 \pm 0.16$		<sup>2</sup> ABELA 83	SPEC	-	$\mu$ X-rays
<sup>1</sup> FETSCHER 84 uses only the measurement of CARR 83.					
<sup>2</sup> Sign of measurement reversed in ABELA 83 to compare with $\mu^+$ measurements.					

## FORM FACTORS FOR RADIATIVE PION AND KAON DECAYS

Updated August 2015 by M. Bychkov (University of Virginia) and G. D’Ambrosio (INFN Sezione di Napoli)

The radiative decays,  $\pi^\pm \rightarrow l^\pm \nu \gamma$  and  $K^\pm \rightarrow l^\pm \nu \gamma$ , with  $l$  standing for an  $e$  or a  $\mu$ , and  $\gamma$  for a real or virtual photon ( $e^+e^-$  pair), provide a powerful tool to investigate the hadronic structure of pions and kaons. The structure-dependent part  $SD_i$  of the amplitude describes the emission of photons from virtual hadronic states, and is parametrized in terms of form factors  $V, A$ , (vector, axial vector), in the standard description [1,2,3,4]. Note that in the Listings below and some literature, equivalent nomenclature  $F_V$  and  $F_A$  for the vector and axial form factors is often used. Exotic, non-standard contributions like  $i = T, S$  (tensor, scalar) have also been considered. Apart from the SD terms, there is also the Inner Bremsstrahlung amplitude, IB, corresponding to photon radiation from external charged particles and described by Low theorem in terms of the physical decay  $\pi^\pm(K^\pm) \rightarrow l^\pm \nu$ . Experiments try to optimize their kinematics so as to minimize the IB part of the amplitude.

The SD amplitude in its standard form is given as

$$M(SD_V) = \frac{-eG_F U_{qq'}}{\sqrt{2}m_P} \epsilon^\mu l^\nu V^P \epsilon_{\mu\nu\sigma\tau} k^\sigma q^\tau \quad (1)$$

$$M(SD_A) = \frac{-ieG_F U_{qq'}}{\sqrt{2}m_P} \epsilon^\mu l^\nu \{A^P[(qk - k^2)g_{\mu\nu} - q_\mu k_\nu] + R^P k^2 g_{\mu\nu}\}, \quad (2)$$

which contains an additional axial form factor  $R^P$  which only can be accessed if the photon remains virtual.  $U_{qq'}$  is the Cabibbo-Kobayashi-Maskawa mixing-matrix element;  $\epsilon^\mu$  is the polarization vector of the photon (or the effective vertex,  $\epsilon^\mu = (e/k^2)\bar{u}(p_-)\gamma^\mu v(p_+)$ , of the  $e^+e^-$  pair);  $\ell^\nu = \bar{u}(p_\nu)\gamma^\nu(1 - \gamma_5)v(p_\ell)$  is the lepton-neutrino current;  $q$  and  $k$  are the meson and photon four-momenta ( $k = p_+ + p_-$  for virtual photons); and  $P$  stands for  $\pi$  or  $K$ .

The pion vector form factor,  $V^\pi$ , is related via CVC (Conserved Vector Current) to the  $\pi^0 \rightarrow \gamma\gamma$  decay width. The constant term is given by  $|V^\pi(0)| = (1/\alpha)\sqrt{2\Gamma_{\pi^0 \rightarrow \gamma\gamma}/\pi m_{\pi^0}}$  [3]. The resulting value,  $V^\pi(0) = 0.0259(9)$ , has been confirmed by calculations based on chiral perturbation theory ( $\chi PT$ ) [4], and by two experiments given in the Listings below. A recent experiment by the PIBETA collaboration [5] obtained a  $V^\pi(0)$  that is in excellent agreement with the CVC hypothesis. It also measured the slope parameter  $a$  in  $V^\pi(s) = V^\pi(0)(1 + a \cdot s)$ , where  $s = (1 - 2E_\gamma/m_\pi)$ , and  $E_\gamma$  is the gamma energy in the pion rest frame:  $a = 0.095 \pm 0.058$ . A functional dependence on  $s$  is expected for all form factors. It becomes non-negligible in the case of  $V^\pi(s)$  when a wide range of photon momenta is recorded; proper treatment in the analysis of  $K$  decays is mandatory.

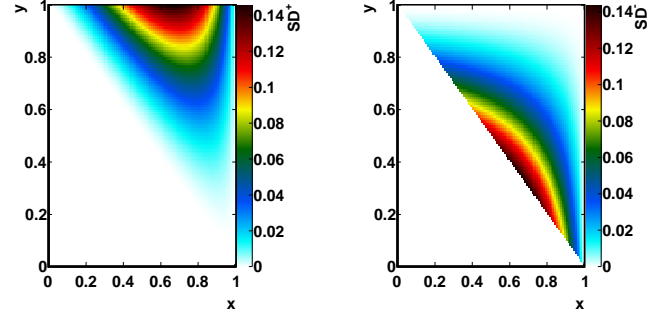
The form factor,  $R^P$ , can be related to the electromagnetic radius,  $r_P$ , of the meson [2]:  $R^P = \frac{1}{3}m_P f_P \langle r_P^2 \rangle$  using PCAC (Partial Conserved Axial vector Current;  $f_P$  is the meson decay constant). In lowest order  $\chi PT$ , the ratio  $A^\pi/V^\pi$  is related to the pion electric polarizability  $\alpha_E = [\alpha/(8\pi^2 m_\pi f_\pi^2)] \times A^\pi/V^\pi$  [6]. The experimental and theoretical status of the pion polarizability is currently an area under discussion and will be addressed in a future edition of this article. The first non-trivial  $\chi PT$  contributions to  $A^K$  and  $V^K$  appear at  $\mathcal{O}(p^4)$  [4], respectively from Gasser-Leutwyler coefficients,  $L_i$ 's, and the anomalous lagrangian:

$$A^K = \frac{4\sqrt{2}M_K}{F_\pi}(L_9^r + L_{10}^r) = 0.042, \quad V^K = \frac{\sqrt{2}M_K}{8\pi^2 F_\pi} = 0.096. \quad (3)$$

$\mathcal{O}(p^6)$  contributions to  $A^K$  can be predicted accurately: they are flat in the momentum dependence and shift the  $\mathcal{O}(p^4)$  value to 0.034.  $\mathcal{O}(p^6)$  contributions to  $V^K$  are model dependent and can be approximated by a form factor linearly dependent on momentum. For example, when looking at the spread of results obtained within two different models, the constant piece of this linear form factor is shifted to  $0.078 \pm 0.005$  [1,2,4].

For decay processes where the photon is real, the partial decay width can be written in analytical form as a sum of IB, SD, and IB/SD interference terms INT [1,4]:

$$\begin{aligned} \frac{d^2\Gamma_{P \rightarrow \ell\nu\gamma}}{dxdy} &= \frac{d^2(\Gamma_{IB} + \Gamma_{SD} + \Gamma_{INT})}{dxdy} \\ &= \frac{\alpha}{2\pi}\Gamma_{P \rightarrow \ell\nu} \frac{1}{(1-r)^2} \left\{ \text{IB}(x, y) \right. \end{aligned}$$



**Figure 1:** Components of the structure dependent terms of the decay width. Left:  $SD^+$ , right:  $SD^-$

$$\begin{aligned} &+ \frac{1}{r} \left( \frac{m_P}{2f_P} \right)^2 \left[ (V+A)^2 SD^+(x, y) + (V-A)^2 SD^-(x, y) \right] \\ &+ \frac{m_P}{f_P} \left[ (V+A) S_{INT}^+(x, y) + (V-A) S_{INT}^-(x, y) \right] \}. \quad (4) \end{aligned}$$

Here

$$\begin{aligned} \text{IB}(x, y) &= \left[ \frac{1-y+r}{x^2(x+y-1-r)} \right] \\ &\quad \left[ x^2 + 2(1-x)(1-r) - \frac{2xr(1-r)}{x+y-1-r} \right] \\ SD^+(x, y) &= (x+y-1-r) \left[ (x+y-1)(1-x) - r \right] \\ SD^-(x, y) &= (1-y+r) \left[ (1-x)(1-y) + r \right] \\ S_{INT}^+(x, y) &= \left[ \frac{1-y+r}{x(x+y-1-r)} \right] \left[ (1-x)(1-x-y) + r \right] \\ S_{INT}^-(x, y) &= \left[ \frac{1-y+r}{x(x+y-1-r)} \right] \left[ x^2 - (1-x)(1-x-y) - r \right] \end{aligned} \quad (5)$$

where  $x = 2E_\gamma/m_P$ ,  $y = 2E_\ell/m_P$ , and  $r = (m_\ell/m_P)^2$ . The structure dependent terms  $SD^+$  and  $SD^-$  are shown in Fig. 1. The  $SD^-$  term is maximized in the same kinematic region where overwhelming IB term dominates (along  $x+y=1$  diagonal). Thus experimental yields with less background are dominated by  $SD^+$  contribution and proportional to  $A^P + V^P$  making simultaneous precise determination of the form factors difficult.

Recently, formulas (4) and (5) have been extended to describe polarized distributions in radiative meson and muon decays [7].

The ‘‘helicity’’ factor  $r$  is responsible for the enhancement of the SD over the IB amplitude in the decays  $\pi^\pm \rightarrow e^\pm \nu \gamma$ , while  $\pi^\pm \rightarrow \mu^\pm \nu \gamma$  is dominated by IB. Interference terms are important for the decay  $K^\pm \rightarrow \mu^\pm \nu \gamma$  [8], but contribute only a few percent correction to pion decays. However, they provide the basis for determining the signs of  $V$  and  $A$ . Radiative corrections to the decay  $\pi^+ \rightarrow e^+ \nu \gamma$  have to be taken into account in the analysis of the precision experiments. They

# Meson Particle Listings

$\pi^\pm$

make up to 4% corrections in the total decay rate [9]. In  $\pi^\pm \rightarrow e^\pm \nu e^\pm e^-$  and  $K^\pm \rightarrow \ell^\pm \nu e^\pm e^-$  decays, all three form factors,  $V^P$ ,  $A^P$ , and  $R^P$ , can be determined [10,11].

We give the experimental  $\pi^\pm$  form factors  $F_T^\pi$ ,  $A_T^\pi$ , and  $R^\pi$  in the Listings below. In the  $K^\pm$  Listings, we give the extracted sum  $A^K + V^K$  and difference  $A^K - V^K$ , as well as  $V^K$ ,  $A^K$  and  $R^K$ . In particular KLOE has measured for the constant piece of the form factor  $A^K + V^K = 0.125 \pm 0.007 \pm 0.001$  [13] while Istra+,  $V^K - A^K = 0.21 \pm 0.04 \pm 0.04$  [14].

Several searches for the exotic form factors  $F_T^K$ ,  $F_T^K$  (tensor), and  $F_S^K$  (scalar) have been pursued in the past. In particular,  $F_T^\pi$  has been brought into focus by experimental as well as theoretical work [12]. New high-statistics data from the PIBETA collaboration have been re-analyzed together with an additional data set optimized for low backgrounds in the radiative pion decay. In particular, lower beam rates have been used in order to reduce the accidental background, thereby making the treatment of systematic uncertainties easier and more reliable. The PIBETA analysis now restricts  $F_T^\pi$  to the range  $-5.2 \times 10^{-4} < F_T^\pi < 4.0 \times 10^{-4}$  at a 90% confidence limit [5]. This result is in excellent agreement with the most recent theoretical work [4].

Precision measurements of radiative pion and kaon decays are effective tools to study QCD in the non-perturbative region and are of interest beyond the scope of radiative decays. Meanwhile other processes such as  $\pi^+ \rightarrow e^+ \nu$  that seem to be better suited to search for new physics at the precision frontier are currently studied. The advantages of such process are the very accurate and reliable theoretical predictions and the more straightforward experimental analysis.

## References

- Phys. Reports **88**, 151 (1982). See our note on “Decay Constants of Charged Pseudoscalar Mesons” elsewhere in this *Review*;  
S.G. Brown and S.A. Bludman, Phys. Rev. **136**, B1160 (1964);  
P. DeBaenst and J. Pestieau, Nuovo Cimento **A53**, 137 (1968).
- W.T. Chu *et al.*, Phys. Rev. **166**, 1577 (1968);  
D.Yu. Bardin and E.A. Ivanov, Sov. J. Part. Nucl. **7**, 286 (1976);  
A. Kersch and F. Scheck, Nucl. Phys. **B263**, 475 (1986).
- V.G. Vaks and B.L. Ioffe, Nuovo Cimento **10**, 342 (1958);  
V.F. Muller, Z. Phys. **173**, 438 (1963).
- C.Q. Geng, I-Lin Ho, and T.H. Wu, Nucl. Phys. **B684**, 281 (2004);  
J. Bijnens and P. Talavera, Nucl. Phys. **B489**, 387 (1997);  
V. Mateu and J. Portoles, Eur. Phys. J. **C52**, 325 (2007);  
R. Unterdorfer, H. Pichl, Eur. Phys. J. **C55**, 273 (2008);  
V. Cirigliano *et al.*, Rev. Mod. Phys. **84**, 399 (2012).
- D. Počanić *et al.*, Phys. Rev. Lett. **93**, 181803 (2004);  
E. Frlež *et al.*, Phys. Rev. Lett. **93**, 181804 (2004);  
M. Bychkov *et al.*, Phys. Rev. Lett. **103**, 051802 (2009).
- J.F. Donoghue and B.R. Holstein, Phys. Rev. **D40**, 2378 (1989).
- E. Gabrielli and L. Trentadue, Nucl. Phys. **B792**, 48 (2008).

- S. Adler *et al.*, Phys. Rev. Lett. **85**, 2256 (2000).
- Yu.M. Bystritsky, E.A. Kuraev, and E.P. Velicheva, Phys. Rev. **D69**, 114004 (2004);  
R. Unterdorfer and H. Pichl have treated radiative corrections of the structure terms to lowest order within  $\chi PT$  for the first time. See the reference under [4].
- S. Egli *et al.*, Phys. Lett. **B175**, 97 (1986).
- A.A. Poblaguev *et al.*, Phys. Rev. Lett. **89**, 061803 (2002).
- A.A. Poblaguev, Phys. Lett. **B238**, 108 (1990);  
V.N. Bolotov *et al.*, Phys. Lett. **B243**, 308 (1990);  
V.M. Belyaev and I.I. Kogan, Phys. Lett. **B280**, 238 (1992);  
A.V. Chernyshev *et al.*, Mod. Phys. Lett. **A12**, 1669 (1997);  
A.A. Poblaguev, Phys. Rev. **D68**, 054020 (2003);  
M.V. Chizhov, Phys. Part. Nucl. Lett. **2**, 193 (2005).
- F. Ambrosino *et al.*, Eur. Phys. J. **C64**, 627 (2009).
- V.A. Duk *et al.*, Phys. Lett. **B695**, 59 (2011).

## $\pi^\pm$ FORM FACTORS

### $F_V$ , VECTOR FORM FACTOR

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0254 ± 0.0017 OUR AVERAGE</b>				
0.0258 ± 0.0017	65k	<sup>1</sup> BYCHKOV	09	PIBE $e^+ \nu \gamma$ at rest
0.014 ± 0.009		<sup>2</sup> BOLOTOV	90B	SPEC 17 GeV $\pi^- \rightarrow e^- \bar{\nu}_e \gamma$
0.023 <sup>+0.015</sup> <sub>-0.013</sub>	98	EGLI	89	SPEC $\pi^+ \rightarrow e^+ \nu_e e^+ e^-$

<sup>1</sup> The BYCHKOV 09  $F_A$  and  $F_V$  results are highly (anti-)correlated:  $F_A + 1.0286 F_V = 0.03853 \pm 0.00014$ .  
<sup>2</sup> BOLOTOV 90B only determines the absolute value.

### $F_A$ , AXIAL-VECTOR FORM FACTOR

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0119 ± 0.0001</b>	65k	<sup>1,2</sup> BYCHKOV	09	PIBE $e^+ \nu \gamma$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0115 ± 0.0004	41k	<sup>1,3</sup> FRLEZ	04	PIBE $\pi^+ \rightarrow e^+ \nu \gamma$ at rest
0.0106 ± 0.0060		<sup>1,4</sup> BOLOTOV	90B	SPEC 17 GeV $\pi^- \rightarrow e^- \bar{\nu}_e \gamma$
0.021 <sup>+0.011</sup> <sub>-0.013</sub>	98	EGLI	89	SPEC $\pi^+ \rightarrow e^+ \nu_e e^+ e^-$
0.0135 ± 0.0016		<sup>1,4</sup> BAY	86	SPEC $\pi^+ \rightarrow e^+ \nu \gamma$
0.006 ± 0.003		<sup>1,4</sup> PIILONEN	86	SPEC $\pi^+ \rightarrow e^+ \nu \gamma$
0.011 ± 0.003		<sup>1,4,5</sup> STETZ	78	SPEC $\pi^+ \rightarrow e^+ \nu \gamma$

<sup>1</sup> These values come from fixing the vector form factor at the CVC prediction,  $F_V = 0.0259 \pm 0.0005$ .  
<sup>2</sup> When  $F_V$  is released, the BYCHKOV 09  $F_A$  is  $0.0117 \pm 0.0017$ , and  $F_A$  and  $F_V$  results are highly (anti-)correlated:  $F_A + 1.0286 F_V = 0.03853 \pm 0.00014$ .

<sup>3</sup> The sign of  $\gamma = F_A / F_V$  is determined to be positive.  
<sup>4</sup> Only the absolute value of  $F_A$  is determined.

<sup>5</sup> The result of STETZ 78 has a two-fold ambiguity. We take the solution compatible with later determinations.

### VECTOR FORM FACTOR SLOPE PARAMETER $a$

This is  $a$  in  $F_V(q^2) = F_V(0) (1 + a q^2)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.10 ± 0.06</b>	65k	BYCHKOV	09	PIBE $e^+ \nu \gamma$ at rest

### $R$ , SECOND AXIAL-VECTOR FORM FACTOR

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.059 <sup>+0.009</sup> <sub>-0.008</sub></b>	98	EGLI	89	SPEC $\pi^+ \rightarrow e^+ \nu_e e^+ e^-$

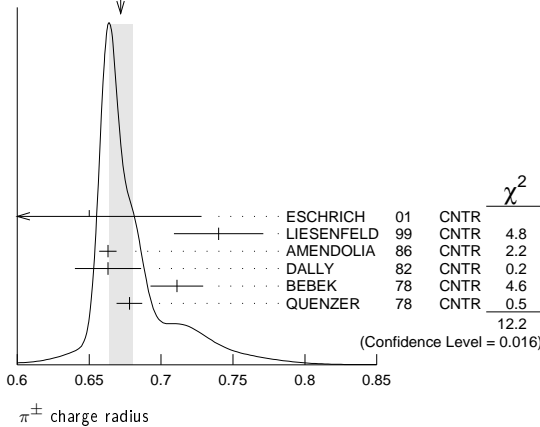
## $\pi^\pm$ CHARGE RADIUS

VALUE (fm)	DOCUMENT ID	TECN	COMMENT
<b>0.672 ± 0.008 OUR AVERAGE</b>	Error includes scale factor of 1.7. See the ideogram below.		
0.65 ± 0.05 ± 0.06	ESCHRICH	01	CNTR $\pi e \rightarrow \pi e$
0.740 ± 0.031	LIESENFELD	99	CNTR $e p \rightarrow e \pi^+ n$
0.663 ± 0.006	AMENDOLIA	86	CNTR $\pi e \rightarrow \pi e$
0.663 ± 0.023	DALLY	82	CNTR $\pi e \rightarrow \pi e$
0.711 ± 0.009 ± 0.016	BEBEK	78	CNTR $e N \rightarrow e \pi N$
0.678 ± 0.004 ± 0.008	QUENZER	78	CNTR $e^+ e^- \rightarrow \pi^+ \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.661 ± 0.012	<sup>1</sup> BIJNENS	98	CNTR $\chi PT$ extraction
0.660 ± 0.024	AMENDOLIA	84	CNTR $\pi e \rightarrow \pi e$
0.78 <sup>+0.09</sup> <sub>-0.10</sub>	ADYLOV	77	CNTR $\pi e \rightarrow \pi e$
0.74 <sup>+0.11</sup> <sub>-0.13</sub>	BARDIN	77	CNTR $e p \rightarrow e \pi^+ n$
0.56 ± 0.04	DALLY	77	CNTR $\pi e \rightarrow \pi e$

See key on page 601

## Meson Particle Listings

$$\pi^\pm, \pi^0$$

<sup>1</sup> BIJNENS 98 fits existing data.WEIGHTED AVERAGE  
0.672±0.008 (Error scaled by 1.7) $\pi^\pm$  REFERENCES

We have omitted some papers that have been superseded by later experiments. The omitted papers may be found in our 1988 edition Physics Letters **B204** 1 (1988).

ADOLPH	15A	PRL 114 062002	C. Adolph <i>et al.</i>	(COMPASS Collab.)
AGUILAR-AR...	15	PRL 115 071801	A.A. Aguilar-Arevalo <i>et al.</i>	(PIENU Collab.)
HOLSTEIN	14	ARNPS 64 51	B. Holstein, S. Scherer	(MASA, MANZ)
BYCHKOV	09	PRL 103 051802	M. Bychkov <i>et al.</i>	(PSI PIBETA Collab.)
FRLEZ	04	PRL 93 181804	E. Frliez <i>et al.</i>	(PSI PIBETA Collab.)
POCANIC	04	PRL 93 181803	D. Pocanic <i>et al.</i>	(PSI PIBETA Collab.)
ESCHRICH	01	PL B522 233	J. Eschrich <i>et al.</i>	(FNAL SELEX Collab.)
LIESENFELD	99	PL B468 20	A. Liesenfeld <i>et al.</i>	
BIJNENS	98	JHEP 9805 014	J. Bijnens <i>et al.</i>	
BRESSI	98	NP B513 555	G. Bressi <i>et al.</i>	
LENZ	98	PL B416 50	S. Lenz <i>et al.</i>	
ASSAMAGAN	96	PR D53 6065	K.A. Assamagan <i>et al.</i>	(PSI, ZURI, VILL+)
KOPTEV	95	JETPL 61 877	V.P. Koptev <i>et al.</i>	(PNPI)
Translated from ZETFP 61 865.				
NUMAO	95	PR D52 4855	T. Numao <i>et al.</i>	(TRIUMF, CRCO)
ASSAMAGAN	94	PL B335 231	K.A. Assamagan <i>et al.</i>	(PSI, ZURI, VILL+)
JECKELMANN	94	PL B335 326	B. Jeckelmann, P.F.A. Goudsmit, H.J. Lelst	(WABRIN+)
CZAPEK	93	PRL 70 17	G. Czappek <i>et al.</i>	(BERN, VILL)
BARANOV	92	SJNP 55 1644	V.A. Baranov <i>et al.</i>	(JINR)
Translated from YAF 55 2940.				
BRITTON	92	PRL 68 3000	D.I. Britton <i>et al.</i>	(TRIUMF, CARL)
Also		PR D49 28	D.I. Britton <i>et al.</i>	(TRIUMF, CARL)
NUMAO	92	MPL A7 3357	T. Numao	(TRIUMF)
BARANOV	91B	SJNP 54 790	V.A. Baranov <i>et al.</i>	(JINR)
Translated from YAF 54 1298.				
DAUM	91	PL B265 425	M. Daum <i>et al.</i>	(VILL)
BOLOTOV	90B	PL B243 308	V.N. Bolotov <i>et al.</i>	(INRM)
EGLI	89	PL B222 533	S. Egli <i>et al.</i>	(SINDRUM Collab.)
Also		PL B175 97	S. Egli <i>et al.</i>	(AACH3, ETH, SIN, ZURI)
PDG	88	PL B204 1	G.P. Yost <i>et al.</i>	(LBL+)
PICCIOTTO	88	PR D37 1131	C.E. Picciotto <i>et al.</i>	(TRIUMF, CNRC)
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RISC, NBS)
KORENCHEN...	87	SJNP 46 192	S.M. Korenchenko <i>et al.</i>	(JINR)
Translated from YAF 46 313.				
AMENDOLIA	86	NP B277 168	S.R. Amendolia <i>et al.</i>	(CERN NA7 Collab.)
BAY	86	PL B174 445	A. Bay <i>et al.</i>	(LAUS, ZURI)
BRYMAN	86	PR D33 1211	D.A. Bryman <i>et al.</i>	(TRIUMF, CNRC)
Also		PRL 50 7	D.A. Bryman <i>et al.</i>	(TRIUMF, CNRC)
JECKELMANN	86B	NP A457 709	B. Jeckelmann <i>et al.</i>	(ETH, FRIB)
Also		PRL 56 1444	B. Jeckelmann <i>et al.</i>	(ETH, FRIB)
PILONEN	86	PRL 57 1402	L.E. Pilonen <i>et al.</i>	(LANL, TEMP, CHIC)
MCFARLANE	85	PR D32 547	W.K. McFarlane <i>et al.</i>	(TEMP, LANL)
ABELA	84	PL 146B 431	R. Abela <i>et al.</i>	(SIN)
Also		PL 74B 126	M. Daum <i>et al.</i>	(SIN)
Also		PR D20 2692	M. Daum <i>et al.</i>	(SIN)
AMENDOLIA	84	PL 146B 116	S.R. Amendolia <i>et al.</i>	(CERN NA7 Collab.)
FETSCHER	84	PL 140B 117	W. Fetscher	(ETH)
ABELA	83	NP A395 413	R. Abela <i>et al.</i>	(BASL, KARL, KARLE)
CARR	83	PRL 51 627	J. Carr <i>et al.</i>	(LBL, NWES, TRIUMF)
COOPER	82	PL 112B 97	A.M. Cooper <i>et al.</i>	(RL)
DALLY	82	PRL 48 375	E.B. Dally <i>et al.</i>	
LU	80	PRL 45 1066	D.C. Lu <i>et al.</i>	(YALE, COLU, JHU)
BEBEK	78	PR D17 1693	C.J. Bebek <i>et al.</i>	
QUENZER	78	PL 76B 512	A. Quenzer <i>et al.</i>	(LALO)
STETZ	78	NP B138 285	A.W. Stetz <i>et al.</i>	(LBL, UCLA)
ADYLOV	77	NP B128 461	G.T. Adylov <i>et al.</i>	
BARDIN	77	NP B120 45	G. Bardin <i>et al.</i>	
DALLY	77	PRL 39 1176	E.B. Dally <i>et al.</i>	
CARTER	76	PRL 37 1380	A.L. Carter <i>et al.</i>	(CARL, CNRC, CHIC+)
KORENCHEN...	76B	JETP 44 35	S.M. Korenchenko <i>et al.</i>	(JINR)
Translated from ZETFP 71 69.				
MARUSHEN...	76	JETPL 23 72	V.I. Marushenko <i>et al.</i>	(PNPI)
Also		Translated from ZETFP 23 80.		
Also		Private Comm.	R.E. Shafer	(FNAL)
Also		Private Comm.	A. Smirnov	(PNPI)
DUNAISEV	73	SJNP 16 292	A.F. Dunaisev <i>et al.</i>	(SERP)
Translated from YAF 16 524.				
AYRES	71	PR D3 1051	D.S. Ayres <i>et al.</i>	(LRL, UCSB)
Also		PR 157 1288	D.S. Ayres <i>et al.</i>	(LRL)
Also		PRL 21 261	D.S. Ayres <i>et al.</i>	(LRL, UCSB)
Also		Thesis UCRL 18369	D.S. Ayres	(LRL)
Also		PR 23 1267	A.J. Greenberg <i>et al.</i>	(LRL, UCSB)
KORENCHEN...	71	SJNP 13 189	S.M. Korenchenko <i>et al.</i>	(JINR)
Translated from YAF 13 339.				

BOOTH	70	PL 32B 723	P.S.L. Booth <i>et al.</i>	(LIVP)
DEPOMMIER	68	NP B4 189	P. Depommier <i>et al.</i>	(CERN)
PETRUKHIN	68	JINR P1 3862	V.I. Petrukhin <i>et al.</i>	(JINR)
HYMAN	67	PL 25B 376	L.G. Hyman <i>et al.</i>	(ANL, CMU, NWES)
NORDBERG	67	PL 24B 594	M.E. Nordberg, F. Lobkowicz, R.L. Burman	(ROCH)
BARDON	66	PRL 16 775	M. Bardon <i>et al.</i>	(COLU)
KINSEY	66	PR 144 1132	K.F. Kinsey, F. Lobkowicz, M.E. Nordberg	(ROCH)
LOBKOWICZ	66	PRL 17 548	F. Lobkowicz <i>et al.</i>	(ROCH, BNL)
BACASTOW	65	PR 139 B407	R.B. Bacastow <i>et al.</i>	(LRL, SLAC)
BERTRAM	65	PR 139 B617	W.K. Bertram <i>et al.</i>	(MICH, CMU)
DUNAISEV	65	JETP 20 58	A.F. Dunaisev <i>et al.</i>	(JINR)
Translated from ZETFP 47 84.				
ECKHAUSE	65	PL 19 348	M. Eckhause <i>et al.</i>	(WILL)
BARTLETT	64	PR 136 B1452	D. Bartlett <i>et al.</i>	(COLU)
DICAPUA	64	PR 133 B1333	M. di Capua <i>et al.</i>	(COLU)
Also		Private Comm.	L. Pondrom	(WISC)
DEPOMMIER	63	PL 5 61	P. Depommier <i>et al.</i>	(CERN)
DEPOMMIER	63B	PL 7 285	P. Depommier <i>et al.</i>	(CERN)
ANDERSON	60	PR 119 2050	H.L. Anderson <i>et al.</i>	(EFI)
CASTAGNOLI	58	PR 112 1779	C. Castagnoli, M. Muchnik	(ROMA)

$$\pi^0$$

$$I^G(J^{PC}) = 1^-(0^+)$$

We have omitted some results that have been superseded by later experiments. The omitted results may be found in our 1988 edition Physics Letters **B204** 1 (1988).

 $\pi^0$  MASS

The value is calculated from  $m_{\pi^\pm}$  and  $(m_{\pi^\pm} - m_{\pi^0})$ . See also the notes under the  $\pi^\pm$  Mass Listings.

VALUE (MeV)	DOCUMENT ID
<b>134.9766±0.0006 OUR FIT</b>	Error includes scale factor of 1.1.

$$m_{\pi^\pm} - m_{\pi^0}$$

Measurements with an error > 0.01 MeV have been omitted.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4.5936 ±0.0005 OUR FIT</b>			
<b>4.5936 ±0.0005 OUR AVERAGE</b>			
4.59364 ±0.00048	CRAWFORD	91	CNTR $\pi^- p \rightarrow \pi^0 n, n$ TOF
4.5930 ±0.0013	CRAWFORD	86	CNTR $\pi^- p \rightarrow \pi^0 n, n$ TOF
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.59366 ±0.00048	CRAWFORD	88B	CNTR See CRAWFORD 91
4.6034 ±0.0052	VASILEVSKY	66	CNTR
4.6056 ±0.0055	CZIRR	63	CNTR

 $\pi^0$  MEAN LIFE

Most experiments measure the  $\pi^0$  width which we convert to a lifetime. ATHERTON 85 is the only direct measurement of the  $\pi^0$  lifetime. Our average based only on indirect measurement yields  $(8.30 \pm 0.19) \times 10^{-17}$  s. The two Primakoff measurements from 1970 have been excluded from our average because they suffered model-related systematics unknown at the time. More information on the  $\pi^0$  lifetime can be found in BERNSTEIN 13.

VALUE ( $10^{-17}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.52±0.18 OUR AVERAGE</b>				Error includes scale factor of 1.2.
8.32±0.15±0.18		<sup>1</sup> LARIN	11	PRMX Primakoff effect
8.5 ±1.1		<sup>2</sup> BYCHKOV	09	PIBE $\pi^+ \rightarrow e^+ \nu \gamma$ at rest
8.4 ±0.5 ±0.5	1182	<sup>3</sup> WILLIAMS	88	CBAL $e^+ e^- \rightarrow e^+ e^- \pi^0$
8.97±0.22±0.17		ATHERTON	85	CNTR Direct measurement
8.2 ±0.4		<sup>4</sup> BROWMAN	74	CNTR Primakoff effect
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.6 ±0.6		BELLETTINI	70	CNTR Primakoff effect
9 ±0.68		KRYSHKIN	70	CNTR Primakoff effect
7.3 ±1.1		BELLETTINI	65B	CNTR Primakoff effect

- <sup>1</sup> LARIN 11 reported  $\Gamma(\pi^0 \rightarrow \gamma\gamma) = 7.82 \pm 0.14 \pm 0.17$  eV which we converted to mean life  $\tau = \hbar/\Gamma(\text{total})$ .
- <sup>2</sup> BYCHKOV 09 obtains this using the conserved-vector-current relation between the vector form factor  $F_V$  and the  $\pi^0$  lifetime.
- <sup>3</sup> WILLIAMS 88 gives  $\Gamma(\gamma\gamma) = 7.7 \pm 0.5 \pm 0.5$  eV. We give here  $\tau = \hbar/\Gamma(\text{total})$ .
- <sup>4</sup> BROWMAN 74 gives a  $\pi^0$  width  $\Gamma = 8.02 \pm 0.42$  eV. The mean life is  $\hbar/\Gamma$ .

 $\pi^0$  DECAY MODES

For decay limits to particles which are not established, see the appropriate Search sections ( $A^0$  (axion) and Other Light Boson ( $X^0$ ) Searches, etc.).

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $2\gamma$	(98.823±0.034) %	S=1.5
$\Gamma_2$ $e^+ e^- \gamma$	( 1.174±0.035) %	S=1.5
$\Gamma_3$ $\gamma$ positronium	( 1.82 ±0.29 ) $\times 10^{-9}$	
$\Gamma_4$ $e^+ e^+ e^- e^-$	( 3.34 ±0.16 ) $\times 10^{-5}$	
$\Gamma_5$ $e^+ e^-$	( 6.46 ±0.33 ) $\times 10^{-8}$	

# Meson Particle Listings

$\pi^0$

$\Gamma_6$	$4\gamma$		$< 2$	$\times 10^{-8}$	CL=90%
$\Gamma_7$	$\nu\bar{\nu}$		$[a] < 2.7$	$\times 10^{-7}$	CL=90%
$\Gamma_8$	$\nu_e\bar{\nu}_e$		$< 1.7$	$\times 10^{-6}$	CL=90%
$\Gamma_9$	$\nu_\mu\bar{\nu}_\mu$		$< 1.6$	$\times 10^{-6}$	CL=90%
$\Gamma_{10}$	$\nu_\tau\bar{\nu}_\tau$		$< 2.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{11}$	$\gamma\nu\bar{\nu}$		$< 6$	$\times 10^{-4}$	CL=90%

## Charge conjugation (C) or Lepton Family number (LF) violating modes

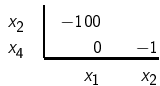
$\Gamma_{12}$	$3\gamma$	C	$< 3.1$	$\times 10^{-8}$	CL=90%
$\Gamma_{13}$	$\mu^+e^-$	LF	$< 3.8$	$\times 10^{-10}$	CL=90%
$\Gamma_{14}$	$\mu^+e^+$	LF	$< 3.4$	$\times 10^{-9}$	CL=90%
$\Gamma_{15}$	$\mu^+e^- + \mu^-e^+$	LF	$< 3.6$	$\times 10^{-10}$	CL=90%

[a] Astrophysical and cosmological arguments give limits of order  $10^{-13}$ ; see the Particle Listings below.

## CONSTRAINED FIT INFORMATION

An overall fit to 2 branching ratios uses 6 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 4.6$  for 4 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.



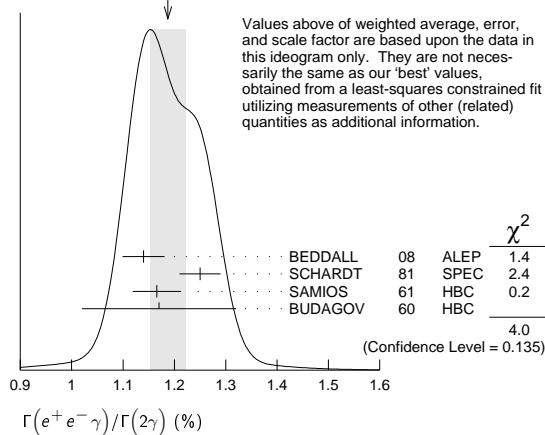
## $\pi^0$ BRANCHING RATIOS

$\Gamma(e^+e^-\gamma)/\Gamma(2\gamma)$	VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_1$
<b>1.188 ± 0.035 OUR FIT</b>	Error				includes scale factor of 1.5.	
<b>1.188 ± 0.034 OUR AVERAGE</b>	Error				includes scale factor of 1.4. See the ideogram below.	
1.140 ± 0.024 ± 0.033	12.5k	5	BEDDALL	08	ALEP $e^+e^- \rightarrow Z \rightarrow \text{hadrons}$	
1.25 ± 0.04			SCHARDT	81	SPEC $\pi^-p \rightarrow n\pi^0$	
1.166 ± 0.047	3071	6	SAMIOS	61	HBC $\pi^-p \rightarrow n\pi^0$	
1.17 ± 0.15	27		BUDAGOV	60	HBC	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1.196			JOSEPH	60	THEO QED calculation	

<sup>5</sup> This BEDDALL 08 value is obtained from ALEPH archived data.

<sup>6</sup> SAMIOS 61 value uses a Panofsky ratio = 1.62.

WEIGHTED AVERAGE  
1.188 ± 0.034 (Error scaled by 1.4)



$\Gamma(\gamma\text{positronium})/\Gamma(2\gamma)$	VALUE (units $10^{-9}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma_1$
<b>1.84 ± 0.29</b>		277	AFANASYEV	90	CNTR $pC$ 70 GeV	

$\Gamma(e^+e^+e^-\gamma)/\Gamma(2\gamma)$	VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma_1$
<b>3.38 ± 0.16 OUR FIT</b>						
<b>3.38 ± 0.16 OUR AVERAGE</b>						
3.46 ± 0.19	30.5k	7	ABOUZAID	08D	KTEV $K_L^0 \rightarrow \pi^0\pi^0\pi_{DD}^0$	
3.18 ± 0.30	146	8	SAMIOS	62B	HBC	

<sup>7</sup> This ABOUZAID 08D value includes all radiative final states. The error includes both statistical and systematic errors. The correlation between the Dalitz-pair planes gives a direct measurement of the  $\pi^0$  parity. The  $\pi^0 2\gamma^*$  form factor is measured and limits are placed on a scalar contribution to the decay.

<sup>8</sup> SAMIOS 62B value uses a Panofsky ratio = 1.62.

## $\Gamma(e^+e^-)/\Gamma_{\text{total}}$

Experimental results are listed; branching ratios corrected for radiative effects are given in the footnotes. BERMAN 60 found  $B(\pi^0 \rightarrow e^+e^-) \geq 4.69 \times 10^{-8}$  via an exact QED calculation.

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>6.46 ± 0.33 OUR AVERAGE</b>					
6.44 ± 0.25 ± 0.22	794	<sup>9</sup> ABOUZAID	07	KTEV	$K_L^0 \rightarrow 3\pi^0$ in flight
6.9 ± 2.3 ± 0.6	21	<sup>10</sup> DESHPANDE	93	SPEC	$K^+ \rightarrow \pi^+\pi^0$
7.6 $\pm_{-2.8}^{+2.9}$ ± 0.5	8	<sup>11</sup> MCFARLAND	93	SPEC	$K_L^0 \rightarrow 3\pi^0$ in flight
• • • We do not use the following data for averages, fits, limits, etc. • • •					
6.09 ± 0.40 ± 0.24	275	<sup>12</sup> ALAVI-HARATI 99c	SPEC	0	Repl. by ABOUZAID 07
<sup>9</sup> ABOUZAID 07 result is for $m_{e^+e^-}/m_{\pi^0} > 0.95$ . With radiative corrections the result becomes $(7.48 \pm 0.29 \pm 0.25) \times 10^{-8}$ .					
<sup>10</sup> The DESHPANDE 93 result with bremsstrahlung radiative corrections is $(8.0 \pm 2.6 \pm 0.6) \times 10^{-8}$ .					
<sup>11</sup> The MCFARLAND 93 result is for $B[\pi^0 \rightarrow e^+e^-, (m_{e^+e^-}/m_{\pi^0})^2 > 0.95]$ . With radiative corrections it becomes $(8.8 \pm_{-3.2}^{+4.5} \pm 0.6) \times 10^{-8}$ .					
<sup>12</sup> ALAVI-HARATI 99c quote result for $B[\pi^0 \rightarrow e^+e^-, (m_{e^+e^-}/m_{\pi^0})^2 > 0.95]$ to minimize radiative contributions from $\pi^0 \rightarrow e^+e^-\gamma$ . After radiative corrections they obtain $(7.04 \pm 0.46 \pm 0.28) \times 10^{-8}$ .					

## $\Gamma(e^+e^-)/\Gamma(2\gamma)$

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.3	90		NIEBUHR	89	SPEC $\pi^-p \rightarrow \pi^0 n$ at rest
<5.3	90		ZEPHAT	87	SPEC $\pi^-p \rightarrow \pi^0 n$ 0.3 GeV/c
1.7 ± 0.6 ± 0.3		59	FRANK	83	SPEC $\pi^-p \rightarrow n\pi^0$
1.8 ± 0.6		58	MISCHKE	82	SPEC See FRANK 83
2.23 $\pm_{-1.10}^{+2.40}$	90	8	FISCHER	78B	SPRK $K^+ \rightarrow \pi^+\pi^0$

## $\Gamma(4\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2</b>	90		MCDONOUGH	88	CBOX $\pi^-p$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<160	90		BOLOTOV	86c	CALO
<440	90	0	AUERBACH	80	CNTR

## $\Gamma(\nu\bar{\nu})/\Gamma_{\text{total}}$

The astrophysical and cosmological limits are many orders of magnitude lower, but we use the best laboratory limit for the Summary Tables.

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.27</b>	90		<sup>13</sup> ARTAMONOV	05A	B949 $K^+ \rightarrow \pi^+\nu\bar{\nu}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.83	90		<sup>13</sup> ATIYA	91	B787 $K^+ \rightarrow \pi^+\nu\nu'$
< 2.9 $\times 10^{-7}$			<sup>14</sup> LAM	91	Cosmological limit
< 3.2 $\times 10^{-7}$			<sup>15</sup> NATALE	91	SN 1987A
< 6.5	90		DORENBOS...	88	CHRM Beam dump, prompt $\nu$
<24	90	0	<sup>13</sup> HERCZEG	81	RVUE $K^+ \rightarrow \pi^+\nu\nu'$
<sup>13</sup> This limit applies to all possible $\nu\nu'$ states as well as to other massless, weakly interacting states.					
<sup>14</sup> LAM 91 considers the production of right-handed neutrinos produced from the cosmic thermal background at the temperature of about the pion mass through the reaction $\gamma\gamma \rightarrow \pi^0 \rightarrow \nu\bar{\nu}$ .					
<sup>15</sup> NATALE 91 considers the excess energy-loss rate from SN 1987A if the process $\gamma\gamma \rightarrow \pi^0 \rightarrow \nu\bar{\nu}$ occurs, permitted if the neutrinos have a right-handed component. As pointed out in LAM 91 (and confirmed by Natale), there is a factor 4 error in the NATALE 91 published result $(0.8 \times 10^{-7})$ .					

## $\Gamma(\nu_e\bar{\nu}_e)/\Gamma_{\text{total}}$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.7</b>	90	DORENBOS..	88	CHRM Beam dump, prompt $\nu$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.1	90	<sup>16</sup> HOFFMAN	88	RVUE Beam dump, prompt $\nu$
<sup>16</sup> HOFFMAN 88 analyzes data from a 400-GeV BEBC beam-dump experiment.				

See key on page 601

## Meson Particle Listings

 $\pi^0, \eta$ 

$\Gamma(\nu_\mu \bar{\nu}_\mu)/\Gamma_{\text{total}}$			$\Gamma_9/\Gamma$		
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.6</b>	90	8.7	AUERBACH 04	LSND	800 MeV $p$ on Cu
<3.1	90		HOFFMAN 88	RVUE	Beam dump, prompt $\nu$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<7.8	90		DORENBOS... 88	CHRM	Beam dump, prompt $\nu$
17 HOFFMAN 88 analyzes data from a 400-GeV BEBC beam-dump experiment.					

$\Gamma(\nu_\tau \bar{\nu}_\tau)/\Gamma_{\text{total}}$			$\Gamma_{10}/\Gamma$		
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.1</b>	90		HOFFMAN 88	RVUE	Beam dump, prompt $\nu$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<4.1	90		DORENBOS... 88	CHRM	Beam dump, prompt $\nu$
18 HOFFMAN 88 analyzes data from a 400-GeV BEBC beam-dump experiment.					

$\Gamma(\gamma\nu)/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$
Standard Model prediction is $6 \times 10^{-18}$ .					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6 \times 10^{-4}$	90	ATIYA 92	CNTR	$K^+ \rightarrow \gamma\nu\pi^+$	

$\Gamma(3\gamma)/\Gamma_{\text{total}}$			$\Gamma_{12}/\Gamma$		
Forbidden by C invariance.					
VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.1</b>	90		MCDONOUGH 88	CBOX	$\pi^- p$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 38	90	0	HIGHLAND 80	CNTR	
<150	90	0	AUERBACH 78	CNTR	
<490	90	0	<sup>19</sup> DUCLOS 65	CNTR	
<490	90		<sup>19</sup> KUTIN 65	CNTR	
<sup>19</sup> These experiments give $B(3\gamma/2\gamma) < 5.0 \times 10^{-6}$ .					

$\Gamma(\mu^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
Forbidden by lepton family number conservation.					
VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.38</b>	90	0	APPEL 00	SPEC	$K^+ \rightarrow \pi^+ \mu^+ e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<16	90		LEE 90	SPEC	$K^+ \rightarrow \pi^+ \mu^+ e^-$
<78	90		CAMPAGNARI 88	SPEC	See LEE 90

$\Gamma(\mu^- e^+)/\Gamma_{\text{total}}$					$\Gamma_{14}/\Gamma$
Forbidden by lepton family number conservation.					
VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.4</b>	90	0	APPEL	00B	B865 $K^+ \rightarrow \pi^+ e^+ \mu^-$

$[\Gamma(\mu^+e^-) + \Gamma(\mu^-e^+)]/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
Forbidden by lepton family number conservation.					
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 0.36	90	ABOUZAID	08c KTEV	$K_L^0 \rightarrow 2\pi^0 \mu^\pm e^\mp$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 17.2	90	KROLAK	94 E799	$\ln K_L^0 \rightarrow 3\pi^0$	
<140		HERCZEG	84 RVUE	$K^+ \rightarrow \pi^+ \mu e$	
< 2 $\times 10^{-6}$		HERCZEG	84 THEO	$\mu^- \rightarrow e^-$ conversion	
< 70	90	BRYMAN	82 RVUE	$K^+ \rightarrow \pi^+ \mu e$	

 $\pi^0$  ELECTROMAGNETIC FORM FACTOR

The amplitude for the process  $\pi^0 \rightarrow e^+ e^- \gamma$  contains a form factor  $F(x)$  at the  $\pi^0 \gamma \gamma$  vertex, where  $x = [m_{e^+ e^-}/m_{\pi^0}]^2$ . The parameter  $a$  in the linear expansion  $F(x) = 1 + ax$  is listed below.

All the measurements except that of BEHREND 91 are in the time-like region of momentum transfer.

LINEAR COEFFICIENT OF  $\pi^0$  ELECTROMAGNETIC FORM FACTOR

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.032 <math>\pm 0.004</math> OUR AVERAGE</b>					
+0.026 $\pm 0.024 \pm 0.048$	7548		FARZANPAY 92	SPEC	$\pi^- p \rightarrow \pi^0 n$ at rest
+0.025 $\pm 0.014 \pm 0.026$	54k		MEIJERDREES 92B	SPEC	$\pi^- p \rightarrow \pi^0 n$ at rest
+0.0326 $\pm 0.0026 \pm 0.0026$	127		20 BEHREND 91	CELL	$e^+ e^- \rightarrow \pi^0$
-0.11 $\pm 0.03 \pm 0.08$	32k		FONVIEILLE 89	SPEC	Radiation corr.
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.12 $\pm 0.05 \pm 0.04$			21 TUPPER 83	THEO	FISCHER 78 data
+0.10 $\pm 0.03$	31k		22 FISCHER 78	SPEC	Radiation corr.
+0.01 $\pm 0.11$	2200		DEVONS 69	OSPK	No radiation corr.
-0.15 $\pm 0.10$	7676		KOBRAK 61	HBC	No radiation corr.
-0.24 $\pm 0.16$	3071		SAMIOS 61	HBC	No radiation corr.

20 BEHREND 91 estimates that their systematic error is of the same order of magnitude as their statistical error, and so we have included a systematic error of this magnitude. The value of  $a$  is obtained by extrapolation from the region of large space-like momentum transfer assuming vector dominance.

21 TUPPER 83 is a theoretical analysis of FISCHER 78 including 2-photon exchange in the corrections.

22 The FISCHER 78 error is statistical only. The result without radiation corrections is  $+0.05 \pm 0.03$ .

 $\pi^0$  REFERENCES

We have omitted some papers that have been superseded by later experiments. The omitted papers may be found in our 1988 edition Physics Letters **B204** 1 (1988).

BERNSTEIN 13	RMP 85 49	A. M. Bernstein, B. R. Holstein	(AMHT, MIT)
LARIN 11	PRL 106 162303	I. Larin <i>et al.</i>	(PrimEx Collab.)
BYCHKOV 09	PRL 103 051802	M. Bychkov <i>et al.</i>	(PSI PIBETA Collab.)
ABOUZAID 08C	PRL 100 131803	E. Abouzaid <i>et al.</i>	(FNAL KTeV Collab.)
ABOUZAID 08D	PRL 100 182001	E. Abouzaid <i>et al.</i>	(FNAL KTeV Collab.)
BEDDALL 08	EPJ C54 365	A. Beddall, A. Beddall	(UGAZ)
ABOUZAID 07	PR D75 012004	E. Abouzaid <i>et al.</i>	(KTeV Collab.)
ARTAMONOV 05A	PR D72 091102	A.V. Artamonov <i>et al.</i>	(BNL E949 Collab.)
AUERBACH 04	PRL 92 091801	L.B. Auerbach <i>et al.</i>	(LSND Collab.)
APPEL 00	PRL 85 2450	R. Appel <i>et al.</i>	(BNL 865 Collab.)
Also Thesis, Yale Univ.			
Also Thesis, Univ. Zurich			
APPEL 00B	PRL 85 2877	R. Appel <i>et al.</i>	(BNL 865 Collab.)
ALAVI-HARATI 99C	PRL 83 922	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
KROLAK 94	PL B320 407	P. Krolak <i>et al.</i>	(EFI, UCLA, COLO, ELMT+)
DESHPANDE 93	PRL 71 27	A. Deshpande <i>et al.</i>	(BNL E851 Collab.)
MC FARLAND 93	PRL 71 31	K.S. McFarland <i>et al.</i>	(EFI, UCLA, COLO+)
ATIYA 92	PRL 69 733	M.S. Atiya <i>et al.</i>	(BNL, LANL, PRIN+)
FARZANPAY 92	PL B278 413	F. Farzanpay <i>et al.</i>	(ORST, TRIU, BRCO+)
MEIJERDREES 92B	PR D45 1439	R. Meijer Drees <i>et al.</i>	(PSI SINDRUM-4 Collab.)
ATIYA 91	PRL 66 2189	M.S. Atiya <i>et al.</i>	(BNL, LANL, PRIN+)
BEHREND 91	ZPHY C49 401	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
CRAWFORD 91	PR D43 46	J.F. Crawford <i>et al.</i>	(VILL, UVA)
LAM 91	PR D44 3345	W.P. Lam, K.W. Ng	(AST)
NATALE 91	PL B258 227	A.A. Natale	(SPIFT)
AFANASYEV 90	PL B236 116	L.G. Afanasyev <i>et al.</i>	(JINR, MOSU, SERP)
Also SJNP 51 664			
Translated from YAF 51 1040.			
LEE 90	PRL 64 1165	A.M. Lee <i>et al.</i>	(BNL, FNAL, VILL, WASH+)
FONVIEILLE 89	PL B233 65	H. Fonvieuille <i>et al.</i>	(CLER, LYON, SACL)
NIEBUHR 89	PR D40 2796	C. Niebuhr <i>et al.</i>	(SINDRUM Collab.)
CAMPAGNARI 88B	PRL 61 2062	C. Campagnari <i>et al.</i>	(BNL, FNAL, PSI+)
CRAWFORD 88B	PL B213 391	J.F. Crawford <i>et al.</i>	(PSI, UVA)
DORENBOSCH 88	ZPHY C40 497	J. Dorenbosch <i>et al.</i>	(CHARM Collab.)
HOFFMAN 88	PL B208 149	C.M. Hoffman	(LANL)
MCDONOUGH 88	PR D38 2121	J.M. McDonough <i>et al.</i>	(TEMP, LANL, CHIC)
PDG 88	PL B204 1	G.P. Yost <i>et al.</i>	(LBL+)
WILLIAMS 88	PR D38 1365	D.A. Williams <i>et al.</i>	(Crystal Ball Collab.)
ZEPHAT 87	JP G13 1375	A.G. Zephath <i>et al.</i>	(OMICRON Collab.)
BOLOTOV 86C	JETPL 43 520	V.N. Bolotov <i>et al.</i>	(INRM)
Translated from ZETFP 43 405.			
CRAWFORD 86	PRL 56 1043	J.F. Crawford <i>et al.</i>	(SIN, UVA)
ATHERTON 85	PL 158B 81	H.W. Atherton <i>et al.</i>	(CERN, ISU, LUND+)
HERCZEG 84	PR D29 1954	P. Herczeg, C.M. Hoffman	(LANL)
FRANK 83	PR D28 423	J.S. Frank <i>et al.</i>	(LANL, ARZS)
TUPPER 83	PR D28 2905	G.B. Tupper, T.R. Grose, M.A. Samuel	(OKSU)
BRYMAN 82	PR D26 2538	D.A. Bryman	(TRIU)
MISCHKE 82	PRL 48 1153	R.E. Mischke <i>et al.</i>	(LANL, ARZS)
HERCZEG 91	PL 100B 347	P. Herczeg, C.M. Hoffman	(LANL)
SCHARDT 81	PR D23 639	M.A. Schardt <i>et al.</i>	(ARZS, LANL)
AUERBACH 80	PL 90B 317	L.B. Auerbach <i>et al.</i>	(TEMP, LASL)
HIGHLAND 80	PRL 44 628	V.L. Highland <i>et al.</i>	(TEMP, LASL)
AUERBACH 78	PRL 41 275	L.B. Auerbach <i>et al.</i>	(TEMP, LASL)
FISCHER 78B	PL 73B 359	J. Fischer <i>et al.</i>	(GEVA, SACL)
FISCHER 78B	PL 73B 364	J. Fischer <i>et al.</i>	(GEVA, SACL)
BROWMAN 74	PL 33 1400	A. Browman <i>et al.</i>	(CORN, BING)
BELLETTINI 70	NC 66A 243	G. Bellettini <i>et al.</i>	(PISA, BON)
KRYSHKIN 70	JETP 30 1037	V.I. Kryshkin, A.G. Sterigov, Y.P. Usov	(TMSK)
Translated from ZETF 57 1917.			
DEVONS 69	PR 184 1356	S. Devons <i>et al.</i>	(COLU, ROMA)
VASILEVSKY 66B	PL 23 281	I.M. Vasilevsky <i>et al.</i>	(JINR)
BELLETTINI 65B	NC 40A 1139	G. Bellettini <i>et al.</i>	(PISA, FIRZ)
DUCLLOS 65	PL 19 253	J. Duclos <i>et al.</i>	(CERN, HEID)
KUTIN 65	JETPL 2 243	V.M. Kuljin, V.I. Petrukhin, Y.D. Prokoshkin	(JINR)
Translated from ZETFP 2 387.			
CZIRR 63	PR 130 341	J.B. Czirr	(LRL)
SAMIOS 62B	PR 126 1844	N.P. Samios <i>et al.</i>	(COLU, BNL)
KOBRAK 61	NC 20 1115	H. Kobrak	(EFI)
SAMIOS 61	PR 121 275	N.P. Samios	(COLU, BNL)
BERMAN 60	NC XVIII 1192	S. Berman, D. Geffen	(JINR)
BUDAGOV 60	JETP 11 755	Y.A. Budagov <i>et al.</i>	(JINR)
Translated from ZETF 38 1047.			
JOSEPH 60	NC 16 997	D.W. Joseph	(EFI)

 $\eta$ 

$$I^G(J^{PC}) = 0^+(0^-+)$$

We have omitted some results that have been superseded by later experiments. The omitted results may be found in our 1988 edition Physics Letters **B204** (1988).

 $\eta$  MASS

Recent measurements resolve the obvious inconsistency in previous  $\eta$  mass measurements in favor of the higher value first reported by NA48 (LA1 02). We use only precise measurements consistent with this higher mass value for our  $\eta$  mass average.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>547.862 <math>\pm 0.017</math> OUR AVERAGE</b>				
547.865 $\pm 0.031 \pm 0.062$		NIKOLAEV 14	CRYB	$\gamma p \rightarrow p \eta$
547.873 $\pm 0.005 \pm 0.027$	1M	GOSLAWSKI 12	SPEC	$d p \rightarrow {}^3\text{He} \eta$
547.874 $\pm 0.007 \pm 0.029$		AMBROSINO 07B	KLOE	$e^+ e^- \rightarrow \phi \rightarrow \eta \gamma$
547.785 $\pm 0.017 \pm 0.057$	16k	MILLER 07	CLEO	$\psi(2S) \rightarrow J/\psi \eta$
547.843 $\pm 0.030 \pm 0.041$	1134	LAI 02	NA48	$\eta \rightarrow 3\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
547.311 $\pm 0.028 \pm 0.032$		1 ABDEL-BARY 05	SPEC	$d p \rightarrow {}^3\text{He} \eta$
547.12 $\pm 0.06 \pm 0.25$		KRUSCH 95D	SPEC	$\gamma p \rightarrow \eta p$ , threshold
547.30 $\pm 0.15$		POUJIN 92	SPEC	$d p \rightarrow {}^3\text{He} \eta$
547.45 $\pm 0.25$		DUANE 74	SPEC	$\pi^- p \rightarrow n$ neutrals

# Meson Particle Listings

## $\eta$

548.2 ± 0.65		FOSTER	65c	HBC
549.0 ± 0.7	148	FOELSCHKE	64	HBC
548.0 ± 1.0	91	ALFF-...	62	HBC
549.0 ± 1.2	53	BASTIEN	62	HBC

<sup>1</sup> ABDEL-BARY 05 disagrees significantly with recent measurements of similar or better precision. See comment in the header.

### $\eta$ WIDTH

This is the partial decay rate  $\Gamma(\eta \rightarrow \gamma\gamma)$  divided by the fitted branching fraction for that mode. See the note at the start of the  $\Gamma(2\gamma)$  data block, next below.

VALUE (keV)	DOCUMENT ID
<b>1.31±0.05 OUR FIT</b>	

### $\eta$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
------	--------------------------------	-----------------------------------

#### Neutral modes

$\Gamma_1$	neutral modes	(72.12±0.34) %	S=1.2
$\Gamma_2$	$2\gamma$	(39.41±0.20) %	S=1.1
$\Gamma_3$	$3\pi^0$	(32.68±0.23) %	S=1.1
$\Gamma_4$	$\pi^0 2\gamma$	( 2.56±0.22 ) × 10 <sup>−4</sup>	
$\Gamma_5$	$2\pi^0 2\gamma$	< 1.2 × 10 <sup>−3</sup>	CL=90%
$\Gamma_6$	$4\gamma$	< 2.8 × 10 <sup>−4</sup>	CL=90%
$\Gamma_7$	invisible	< 1.0 × 10 <sup>−4</sup>	CL=90%

#### Charged modes

$\Gamma_8$	charged modes	(28.10±0.34) %	S=1.2
$\Gamma_9$	$\pi^+ \pi^- \pi^0$	(22.92±0.28) %	S=1.2
$\Gamma_{10}$	$\pi^+ \pi^- \gamma$	( 4.22±0.08 ) %	S=1.1
$\Gamma_{11}$	$e^+ e^- \gamma$	( 6.9 ± 0.4 ) × 10 <sup>−3</sup>	S=1.3
$\Gamma_{12}$	$\mu^+ \mu^- \gamma$	( 3.1 ± 0.4 ) × 10 <sup>−4</sup>	
$\Gamma_{13}$	$e^+ e^-$	< 2.3 × 10 <sup>−6</sup>	CL=90%
$\Gamma_{14}$	$\mu^+ \mu^-$	( 5.8 ± 0.8 ) × 10 <sup>−6</sup>	
$\Gamma_{15}$	$2e^+ 2e^-$	( 2.40±0.22 ) × 10 <sup>−5</sup>	
$\Gamma_{16}$	$\pi^+ \pi^- e^+ e^- (\gamma)$	( 2.68±0.11 ) × 10 <sup>−4</sup>	
$\Gamma_{17}$	$e^+ e^- \mu^+ \mu^-$	< 1.6 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{18}$	$2\mu^+ 2\mu^-$	< 3.6 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{19}$	$\mu^+ \mu^- \pi^+ \pi^-$	< 3.6 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{20}$	$\pi^+ e^- \overline{\nu}_e + \text{c.c.}$	< 1.7 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{21}$	$\pi^+ \pi^- 2\gamma$	< 2.1 × 10 <sup>−3</sup>	
$\Gamma_{22}$	$\pi^+ \pi^- \pi^0 \gamma$	< 5 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{23}$	$\pi^0 \mu^+ \mu^- \gamma$	< 3 × 10 <sup>−6</sup>	CL=90%

### Charge conjugation (C), Parity (P), Charge conjugation × Parity (CP), or Lepton Family number (LF) violating modes

$\Gamma_{24}$	$\pi^0 \gamma$	C	< 9 × 10 <sup>−5</sup>	CL=90%
$\Gamma_{25}$	$\pi^+ \pi^-$	P,CP	< 1.3 × 10 <sup>−5</sup>	CL=90%
$\Gamma_{26}$	$2\pi^0$	P,CP	< 3.5 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{27}$	$2\pi^0 \gamma$	C	< 5 × 10 <sup>−4</sup>	CL=90%
$\Gamma_{28}$	$3\pi^0 \gamma$	C	< 6 × 10 <sup>−5</sup>	CL=90%
$\Gamma_{29}$	$3\gamma$	C	< 1.6 × 10 <sup>−5</sup>	CL=90%
$\Gamma_{30}$	$4\pi^0$	P,CP	< 6.9 × 10 <sup>−7</sup>	CL=90%
$\Gamma_{31}$	$\pi^0 e^+ e^-$	C	[a] < 4 × 10 <sup>−5</sup>	CL=90%
$\Gamma_{32}$	$\pi^0 \mu^+ \mu^-$	C	[a] < 5 × 10 <sup>−6</sup>	CL=90%
$\Gamma_{33}$	$\mu^+ e^- + \mu^- e^+$	LF	< 6 × 10 <sup>−6</sup>	CL=90%

[a] C parity forbids this to occur as a single-photon process.

### CONSTRAINED FIT INFORMATION

An overall fit to 2 decay rate and 19 branching ratios uses 50 measurements and one constraint to determine 9 parameters. The overall fit has a  $\chi^2 = 43.8$  for 42 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	24							
$x_4$	4	1						
$x_9$	−73	−80	−4					
$x_{10}$	−56	−60	−3	61				
$x_{11}$	−5	−5	0	−6	−4			
$x_{12}$	−1	0	0	−1	0	0		
$x_{16}$	0	0	0	0	0	0	0	
$\Gamma$	−14	−3	−32	11	8	1	0	0
	$x_2$	$x_3$	$x_4$	$x_9$	$x_{10}$	$x_{11}$	$x_{12}$	$x_{16}$

Mode	Rate (keV)	Scale factor
$\Gamma_2$ $2\gamma$	0.515 ± 0.018	
$\Gamma_3$ $3\pi^0$	0.427 ± 0.015	
$\Gamma_4$ $\pi^0 2\gamma$	(3.34 ± 0.28 ) × 10 <sup>−4</sup>	
$\Gamma_9$ $\pi^+ \pi^- \pi^0$	0.299 ± 0.011	
$\Gamma_{10}$ $\pi^+ \pi^- \gamma$	0.0551 ± 0.0022	
$\Gamma_{11}$ $e^+ e^- \gamma$	0.0090 ± 0.0006	1.2
$\Gamma_{12}$ $\mu^+ \mu^- \gamma$	(4.1 ± 0.5 ) × 10 <sup>−4</sup>	
$\Gamma_{16}$ $\pi^+ \pi^- e^+ e^- (\gamma)$	(3.50 ± 0.19 ) × 10 <sup>−4</sup>	

### $\eta$ DECAY RATES

#### $\Gamma(2\gamma)$

#### $\Gamma_2$

See the table immediately above giving the fitted decay rates. Following the advice of NEFKENS 02, we have removed the Primakoff-effect measurement from the average. See also the “Note on the Decay Width  $\Gamma(\eta \rightarrow \gamma\gamma)$ ,” in our 1994 edition, Phys. Rev. **D50**, 1 August 1994, Part I, p. 1451, for a discussion of the various measurements.

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.515±0.018 OUR FIT</b>				
<b>0.516±0.018 OUR AVERAGE</b>				
0.520±0.020±0.013		BABUSCI	13A	KLOE $e^+ e^- \rightarrow e^+ e^- \eta$
0.51 ± 0.12 ± 0.05	36	BARU	90	MD1 $e^+ e^- \rightarrow e^+ e^- \eta$
0.490±0.010±0.048	2287	ROE	90	ASP $e^+ e^- \rightarrow e^+ e^- \eta$
0.514 ± 0.017 ± 0.035	1295	WILLIAMS	88	CBAL $e^+ e^- \rightarrow e^+ e^- \eta$
0.53 ± 0.04 ± 0.04		BARTEL	85E	JADE $e^+ e^- \rightarrow e^+ e^- \eta$
0.476 ± 0.062		<sup>1</sup> RODRIGUES	08	CNTR Reanalysis
0.64 ± 0.14 ± 0.13		AIHARA	86	TPC $e^+ e^- \rightarrow e^+ e^- \eta$
0.56 ± 0.16	56	WEINSTEIN	83	CBAL $e^+ e^- \rightarrow e^+ e^- \eta$
0.324±0.046		BROWMAN	74B	CNTR Primakoff effect
1.00 ± 0.22		<sup>2</sup> BEMPORAD	67	CNTR Primakoff effect

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> RODRIGUES 08 uses a more sophisticated calculation for the inelastic background due to incoherent photoproduction to reanalyze the  $\eta$  photoproduction data on Be and Cu at 9 GeV from BROWMAN 74B. This brings the value of  $\Gamma(\eta \rightarrow 2\gamma)$  in line with direct measurements of the width. The error here is only statistical.

<sup>2</sup> BEMPORAD 67 gives  $\Gamma(2\gamma) = 1.21 \pm 0.26$  keV assuming  $\Gamma(2\gamma)/\Gamma(\text{total}) = 0.314$ . Bemporad private communication gives  $\Gamma(2\gamma)^2/\Gamma(\text{total}) = 0.380 \pm 0.083$ . We evaluate this using  $\Gamma(2\gamma)/\Gamma(\text{total}) = 0.38 \pm 0.01$ . Not included in average because the uncertainty resulting from the separation of the coulomb and nuclear amplitudes has apparently been underestimated.

#### $\Gamma(\pi^0 2\gamma)$

#### $\Gamma_4$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.334±0.028 OUR FIT</b>				
<b>0.33 ± 0.03</b>	1200	NEFKENS	14	CRYB $\gamma p \rightarrow \eta p$

### $\eta$ BRANCHING RATIOS

#### Neutral modes

#### $\Gamma(\text{neutral modes})/\Gamma_{\text{total}}$

#### $\Gamma_1/\Gamma = (\Gamma_2+\Gamma_3+\Gamma_4)/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.7212±0.0034 OUR FIT</b>				Error includes scale factor of 1.2.
<b>0.705 ± 0.008</b>	16k	BASILE	71D	CNTR MM spectrometer
0.79 ± 0.08		BUNIATOV	67	OSPK

#### $\Gamma(2\gamma)/\Gamma_{\text{total}}$

#### $\Gamma_2/\Gamma$

VALUE (units 10 <sup>−2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>39.41±0.20 OUR FIT</b>				Error includes scale factor of 1.1.
<b>39.49±0.17±0.30</b>	65k	ABEGG	96	SPEC $p d \rightarrow {}^3\text{He} \eta$
38.45 ± 0.40 ± 0.36	14k	<sup>1</sup> LOPEZ	07	CLEO $\psi(2S) \rightarrow J/\psi \eta$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> Not independent of other results listed for LOPEZ 07. Assuming decays of  $\eta \rightarrow \gamma\gamma$ ,  $3\pi^0$ ,  $\pi^+ \pi^- \pi^0$ ,  $\pi^+ \pi^- \gamma$ , and  $e^+ e^- \gamma$  account for all  $\eta$  decays within a contribution of 0.3% to the systematic error.



See key on page 601

## Meson Particle Listings

 $\eta$ 

$\Gamma(2\gamma)/\Gamma(\text{neutral modes})$		$\Gamma_2/\Gamma_1 = \Gamma_2/(\Gamma_2+\Gamma_3+\Gamma_4)$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.5465 \pm 0.0019</math> OUR FIT</b>				
<b><math>0.548 \pm 0.023</math> OUR AVERAGE</b>				Error includes scale factor of 1.5.
$0.535 \pm 0.018$		BUTTRAM	70	OSPK
$0.59 \pm 0.033$		BUNIATOV	67	OSPK
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.52 \pm 0.09$	88	ABROSIMOV	80	HLBC
$0.60 \pm 0.14$	113	KENDALL	74	OSPK
$0.57 \pm 0.09$		STRUGALSKI	71	HLBC
$0.579 \pm 0.052$		FELDMAN	67	OSPK
$0.416 \pm 0.044$		DIGIUGNO	66	CNTR Error doubled
$0.44 \pm 0.07$		GRUNHAUS	66	OSPK
$0.39 \pm 0.06$		<sup>1</sup> JONES	66	CNTR

<sup>1</sup> This result from combining cross sections from two different experiments.

$\Gamma(3\pi^0)/\Gamma_{\text{total}}$		$\Gamma_3/\Gamma$		
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>32.68 \pm 0.23</math> OUR FIT</b>				Error includes scale factor of 1.1.
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$34.03 \pm 0.56 \pm 0.49$	1821	<sup>1</sup> LOPEZ	07	CLEO $\psi(2S) \rightarrow J/\psi \eta$
<sup>1</sup> Not independent of other results listed for LOPEZ 07. Assuming decays of $\eta \rightarrow \gamma\gamma$ , $3\pi^0$ , $\pi^+\pi^-\pi^0$ , $\pi^+\pi^-\gamma$ , and $e^+e^-\gamma$ account for all $\eta$ decays within a contribution of 0.3% to the systematic error.				

$\Gamma(3\pi^0)/\Gamma(\text{neutral modes})$		$\Gamma_3/\Gamma_1 = \Gamma_3/(\Gamma_2+\Gamma_3+\Gamma_4)$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.4531 \pm 0.0019</math> OUR FIT</b>				
<b><math>0.439 \pm 0.024</math></b>		BUTTRAM	70	OSPK
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.44 \pm 0.08$	75	ABROSIMOV	80	HLBC
$0.32 \pm 0.09$		STRUGALSKI	71	HLBC
$0.41 \pm 0.033$		BUNIATOV	67	OSPK Not indep. of $\Gamma(2\gamma)/\Gamma(\text{neutral modes})$
$0.177 \pm 0.035$		FELDMAN	67	OSPK
$0.209 \pm 0.054$		DIGIUGNO	66	CNTR Error doubled
$0.29 \pm 0.10$		GRUNHAUS	66	OSPK

$\Gamma(3\pi^0)/\Gamma(2\gamma)$		$\Gamma_3/\Gamma_2$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.829 \pm 0.006</math> OUR FIT</b>				
<b><math>0.829 \pm 0.007</math> OUR AVERAGE</b>				
$0.884 \pm 0.022 \pm 0.019$	1821	LOPEZ	07	CLEO $\psi(2S) \rightarrow J/\psi \eta$
$0.817 \pm 0.012 \pm 0.032$	17.4k	<sup>1</sup> AKHMETSHIN	05	CMD2 $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
$0.826 \pm 0.024$		ACHASOV	00D	SND $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
$0.832 \pm 0.005 \pm 0.012$		KRUSCHE	95D	SPEC $\gamma p \rightarrow \eta p$ , threshold
$0.841 \pm 0.034$		AMSLER	93	CBAR $\bar{p}p \rightarrow \pi^+\pi^-\eta$ at rest
$0.822 \pm 0.009$		ALDE	84	GAM2
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.796 \pm 0.016 \pm 0.016$		ACHASOV	00	SND See ACHASOV 00D
$0.91 \pm 0.14$		COX	70B	HBC
$0.75 \pm 0.09$		DEVONS	70	OSPK
$0.88 \pm 0.16$		BALTAY	67D	DBC
$1.1 \pm 0.2$		CENCE	67	OSPK
$1.25 \pm 0.39$		BACCI	63	CNTR Inverse BR reported

<sup>1</sup> Uses result from AKHMETSHIN 01B.

$\Gamma(\pi^0 2\gamma)/\Gamma_{\text{total}}$		$\Gamma_4/\Gamma$		
VALUE (units $10^{-4}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.56 \pm 0.22</math> OUR FIT</b>				Early results are summarized in the review by LANDSBERG 85.
<b><math>2.21 \pm 0.24 \pm 0.47</math></b>	$\approx 500$	<sup>1</sup> PRAKHOV	08	CRYB $\pi^- p \rightarrow \eta n \approx$ threshold
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.5 \pm 0.7 \pm 0.6$	1.6k	<sup>2,3</sup> PRAKHOV	05	CRYB See PRAKHOV 08
$< 8.4$	90	ACHASOV	01D	SND $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
$< 30$	90	DAVYDOV	81	GAM2 $\pi^- p \rightarrow \eta n$

<sup>1</sup> PRAKHOV 08 is a reanalysis of the data of PRAKHOV 05, using for the first time the invariant-mass spectrum of the two photons.<sup>2</sup> Normalized using  $\Gamma(\eta \rightarrow 2\gamma)/\Gamma = 0.3943 \pm 0.0026$ .<sup>3</sup> This measurement and the independent analysis of the same data by KNECHT 04 both imply a lower value of  $\Gamma(\pi^0 2\gamma)$  than the one obtained by ALDE 84 from  $\Gamma(\pi^0 2\gamma)/\Gamma(2\gamma)$ .

$\Gamma(\pi^0 2\gamma)/\Gamma(2\gamma)$		$\Gamma_4/\Gamma_2$		
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG COMMENT
<b><math>0.65 \pm 0.06</math> OUR FIT</b>				
<b><math>1.8 \pm 0.4</math></b>		ALDE	84	GAM2 0
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.5 \pm 0.6$	70	BINON	82	GAM2 See ALDE 84

$\Gamma(\pi^0 2\gamma)/\Gamma(3\pi^0)$		$\Gamma_4/\Gamma_3$		
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>7.8 \pm 0.7</math> OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.3 \pm 2.8 \pm 1.4$		<sup>1</sup> KNECHT	04	CRYB $\pi^- p \rightarrow n\eta$
<sup>1</sup> Independent analysis of same data as PRAKHOV 05.				

$\Gamma(2\pi^0 2\gamma)/\Gamma_{\text{total}}$		$\Gamma_5/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.2 \times 10^{-3}</math></b>	90	<sup>1</sup> NEFKENS	05A	CRYB $p(720 \text{ MeV}/c) \pi^- \rightarrow n\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.0 \times 10^{-3}$	90	BLIK	07	GAM4 $\pi^- p \rightarrow \eta n$
<sup>1</sup> Measurement is done in limited $\gamma\gamma$ energy range.				

$\Gamma(4\gamma)/\Gamma_{\text{total}}$		$\Gamma_6/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.8 \times 10^{-4}</math></b>	90	BLIK	07	GAM4 $\pi^- p \rightarrow \eta n$

$\Gamma(\text{invisible})/\Gamma(2\gamma)$		$\Gamma_7/\Gamma_2$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.6 \times 10^{-4}</math></b>	90	<sup>1</sup> ABLIKIM	13	BES3 $J/\psi \rightarrow \phi \eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.65 \times 10^{-3}$	90	<sup>2</sup> ABLIKIM	06Q	BES2 $J/\psi \rightarrow \phi \eta$
<sup>1</sup> Based on 225M $J/\psi$ decays.				
<sup>2</sup> Based on 58M $J/\psi$ decays.				

## Charged modes

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$		$\Gamma_9/\Gamma$		
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>22.92 \pm 0.28</math> OUR FIT</b>				Error includes scale factor of 1.2.
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$22.60 \pm 0.35 \pm 0.29$	3915	<sup>1</sup> LOPEZ	07	CLEO $\psi(2S) \rightarrow J/\psi \eta$
<sup>1</sup> Not independent of other results listed for LOPEZ 07. Assuming decays of $\eta \rightarrow \gamma\gamma$ , $3\pi^0$ , $\pi^+\pi^-\pi^0$ , $\pi^+\pi^-\gamma$ , and $e^+e^-\gamma$ account for all $\eta$ decays within a contribution of 0.3% to the systematic error.				

$\Gamma(\text{neutral modes})/\Gamma(\pi^+\pi^-\pi^0)$		$\Gamma_1/\Gamma_9 = (\Gamma_2+\Gamma_3+\Gamma_4)/\Gamma_9$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.15 \pm 0.05</math> OUR FIT</b>				Error includes scale factor of 1.2.
<b><math>3.26 \pm 0.30</math> OUR AVERAGE</b>				
$2.54 \pm 1.89$	74	KENDALL	74	OSPK
$3.4 \pm 1.1$	29	AGUILAR...	72B	HBC
$2.83 \pm 0.80$	70	<sup>1</sup> BLOODW...	72B	HBC
$3.6 \pm 0.6$	244	FLATTE	67B	HBC
$2.89 \pm 0.56$		ALFF...	66	HBC
$3.6 \pm 0.8$	50	KRAEMER	64	DBC
$3.8 \pm 1.1$		PAULI	64	DBC
<sup>1</sup> Error increased from published value 0.5 by Bloodworth (private communication).				

$\Gamma(2\gamma)/\Gamma(\pi^+\pi^-\pi^0)$		$\Gamma_2/\Gamma_9$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.720 \pm 0.028</math> OUR FIT</b>				Error includes scale factor of 1.2.
<b><math>1.70 \pm 0.04</math> OUR AVERAGE</b>				
$1.704 \pm 0.032 \pm 0.026$	3915	<sup>1</sup> LOPEZ	07	CLEO $\psi(2S) \rightarrow J/\psi \eta$
$1.61 \pm 0.14$		ABLIKIM	06E	BES2 $e^+e^- \rightarrow J/\psi \rightarrow \eta\gamma$
$1.78 \pm 0.10 \pm 0.13$	1077	AMSLER	95	CBAR $\bar{p}p \rightarrow \pi^+\pi^-\eta$ at rest
$1.72 \pm 0.25$	401	BAGLIN	69	HLBC
$1.61 \pm 0.39$		FOSTER	65	HBC
<sup>1</sup> LOPEZ 07 reports $\Gamma(\eta \rightarrow \pi^+\pi^-\pi^0) / \Gamma(\eta \rightarrow 2\gamma) = \Gamma_9/\Gamma_2 = 0.587 \pm 0.011 \pm 0.009$ .				

$\Gamma(3\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$		$\Gamma_3/\Gamma_9$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.426 \pm 0.026</math> OUR FIT</b>				Error includes scale factor of 1.2.
<b><math>1.48 \pm 0.05</math> OUR AVERAGE</b>				
$1.46 \pm 0.03 \pm 0.09$		ACHASOV	06A	SND $e^+e^- \rightarrow \eta\gamma$
$1.52 \pm 0.04 \pm 0.08$	23k	<sup>1</sup> AKHMETSHIN	01B	CMD2 $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
$1.44 \pm 0.09 \pm 0.10$	1627	AMSLER	95	CBAR $\bar{p}p \rightarrow \pi^+\pi^-\eta$ at rest
$1.50 \pm 0.15 \pm 0.29$	199	BAGLIN	69	HLBC
$1.47 \pm 0.20 \pm 0.17$		BULLOCK	68	HLBC
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.3 \pm 0.4$		BAGLIN	67B	HLBC
$0.90 \pm 0.24$		FOSTER	65	HBC
$2.0 \pm 1.0$		FOELSCHE	64	HBC
$0.83 \pm 0.32$		CRAWFORD	63	HBC
<sup>1</sup> AKHMETSHIN 01B uses results from AKHMETSHIN 99F.				

$\Gamma(\pi^+\pi^-\pi^0)/[\Gamma(2\gamma) + \Gamma(3\pi^0)]$		$\Gamma_9/(\Gamma_2+\Gamma_3)$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.318 \pm 0.005</math> OUR FIT</b>				Error includes scale factor of 1.2.
<b><math>0.304 \pm 0.012</math></b>		ACHASOV	00D	SND $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.3141 \pm 0.0081 \pm 0.0058$		ACHASOV	00B	SND See ACHASOV 00D

# Meson Particle Listings

$\eta$

## $\Gamma(\pi^+\pi^-\gamma)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.22±0.08 OUR FIT</b>	Error includes scale factor of 1.1.			

• • • We do not use the following data for averages, fits, limits, etc. • • •  
3.96±0.14±0.14 859 <sup>1</sup> LOPEZ 07 CLEO  $\psi(2S) \rightarrow J/\psi\eta$

<sup>1</sup> Not independent of other results listed for LOPEZ 07. Assuming decays of  $\eta \rightarrow \gamma\gamma$ ,  $3\pi^0$ ,  $\pi^+\pi^-\pi^0$ ,  $\pi^+\pi^-\gamma$ , and  $e^+e^-\gamma$  account for all  $\eta$  decays within a contribution of 0.3% to the systematic error.

## $\Gamma(\pi^+\pi^-\gamma)/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_{10}/\Gamma_9$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1842±0.0027 OUR FIT</b>				
<b>0.1847±0.0030 OUR AVERAGE</b>	Error includes scale factor of 1.1.			

0.1856±0.0005±0.0028 200k BABUSCI 13 KLOE  $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$   
0.175 ±0.007 ±0.006 859 LOPEZ 07 CLEO  $\psi(2S) \rightarrow J/\psi\eta$   
• • • We do not use the following data for averages, fits, limits, etc. • • •

0.209 ±0.004 18k THALER 73 ASPK  
0.201 ±0.006 7250 GORMLEY 70 ASPK  
0.28 ±0.04 BALTAY 67b DBC  
0.25 ±0.035 LITCHFIELD 67 DBC  
0.30 ±0.06 CRAWFORD 66 HBC  
0.196 ±0.041 FOSTER 65c HBC

## $\Gamma(e^+e^-\gamma)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.9 ±0.4 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>6.7 ±0.5 OUR AVERAGE</b>	Error includes scale factor of 1.2.			

6.6 ±0.4 ±0.4 1345 BERGHAUSER 11 SPEC  $\gamma p \rightarrow p\eta$   
7.8 ±0.5 ±0.8 435 ±31 BERLOWSKI 08 WASA  $p d \rightarrow {}^3\text{He} \eta$   
5.15±0.62±0.74 283 ACHASOV 01b SND  $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$   
7.10±0.64±0.46 323 AKHMETSHIN 01 CMD2  $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$   
• • • We do not use the following data for averages, fits, limits, etc. • • •  
9.4 ±0.7 ±0.5 172 <sup>1</sup> LOPEZ 07 CLEO  $\psi(2S) \rightarrow J/\psi\eta$

<sup>1</sup> Not independent of other results listed for LOPEZ 07. Assuming decays of  $\eta \rightarrow \gamma\gamma$ ,  $3\pi^0$ ,  $\pi^+\pi^-\pi^0$ ,  $\pi^+\pi^-\gamma$ , and  $e^+e^-\gamma$  account for all  $\eta$  decays within a contribution of 0.3% to the systematic error.

## $\Gamma(e^+e^-\gamma)/\Gamma(\pi^+\pi^-\gamma)$ $\Gamma_{11}/\Gamma_{10}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.163±0.011 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.237±0.021±0.015</b>	172	LOPEZ 07	CLEO	$\psi(2S) \rightarrow J/\psi\eta$

## $\Gamma(e^+e^-\gamma)/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_{11}/\Gamma_9$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.00±0.19 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>2.1 ±0.5</b>	80	JANE 75b	OSPK	See the erratum

## $\Gamma(\text{neutral modes})/[\Gamma(\pi^+\pi^-\pi^0) + \Gamma(\pi^+\pi^-\gamma) + \Gamma(e^+e^-\gamma)]$ $\Gamma_1/(\Gamma_9+\Gamma_{10}+\Gamma_{11}) = (\Gamma_2+\Gamma_3+\Gamma_4)/(\Gamma_9+\Gamma_{10}+\Gamma_{11})$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.59±0.04 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>2.64±0.23</b>		BALTAY 67b	DBC	

• • • We do not use the following data for averages, fits, limits, etc. • • •  
4.5 ±1.0 280 <sup>1</sup> JAMES 66 HBC  
3.20±1.26 53 <sup>1</sup> BASTIEN 62 HBC  
2.5 ±1.0 10 <sup>1</sup> PICKUP 62 HBC

<sup>1</sup> These experiments are not used in the averages as they do not separate clearly  $\eta \rightarrow \pi^+\pi^-\pi^0$  and  $\eta \rightarrow \pi^+\pi^-\gamma$  from each other. The reported values thus probably contain some unknown fraction of  $\eta \rightarrow \pi^+\pi^-\gamma$ .

## $\Gamma(2\gamma)/[\Gamma(\pi^+\pi^-\pi^0) + \Gamma(\pi^+\pi^-\gamma) + \Gamma(e^+e^-\gamma)]$ $\Gamma_2/(\Gamma_9+\Gamma_{10}+\Gamma_{11})$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.417±0.023 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>1.1 ±0.4 OUR AVERAGE</b>				

1.51 ±0.93 75 KENDALL 74 OSPK  
0.99 ±0.48 CRAWFORD 63 HBC

## $\Gamma(\mu^+\mu^-\gamma)/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.1±0.4 OUR FIT</b>				
<b>3.1±0.4</b>	600	DZHELYADIN 80	SPEC	$\pi^- p \rightarrow \eta n$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
1.5 ±0.75 100 BUSHNIN 78 SPEC See DZHELYADIN 80

## $\Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_{13}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.3 × 10<sup>-6</sup></b>	90	AGAKISHIEV 14		$pp \rightarrow \eta + X$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<5.6 × 10<sup>-6</sup> 90 <sup>1</sup> AGAKISHIEV 12a SPEC  $pp \rightarrow \eta + X$   
<2.7 × 10<sup>-5</sup> 90 BERLOWSKI 08 WASA  $p d \rightarrow {}^3\text{He} \eta$   
<0.77 × 10<sup>-4</sup> 90 BROWDER 97b CLE2  $e^+e^- \simeq 10.5 \text{ GeV}$   
<2 × 10<sup>-4</sup> 90 WHITE 96 SPEC  $p d \rightarrow \eta^3\text{He}$   
<3 × 10<sup>-4</sup> 90 DAVIES 74 RVUE USES ESTEN 67

<sup>1</sup> AGAKISHIEV 12a uses a data sample of 3.5 GeV proton beam collisions on liquid hydrogen target collected by the HADES detector.

## $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$ $\Gamma_{14}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.8±0.8 OUR AVERAGE</b>					

5.7±0.7±0.5 114 ABEGG 94 SPEC  $p d \rightarrow \eta^3\text{He}$   
6.5±2.1 27 DZHELYADIN 80b SPEC  $\pi^- p \rightarrow \eta n$   
• • • We do not use the following data for averages, fits, limits, etc. • • •

5.6+0.6-0.7±0.5 100 KESSLER 93 SPEC See ABEGG 94  
< 20 95 0 WEHMANN 68 OSPK

## $\Gamma(\mu^+\mu^-)/\Gamma(2\gamma)$ $\Gamma_{14}/\Gamma_2$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5.9±2.2	HYAMS 69	OSPK	

## $\Gamma(2e^+2e^-)/\Gamma_{\text{total}}$ $\Gamma_{15}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
2.4±0.2±0.1	362	<sup>1</sup> AMBROSINO 11b	KLOE	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$	

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<9.7 90 BERLOWSKI 08 WASA  $p d \rightarrow {}^3\text{He} \eta$   
<6.9 90 AKHMETSHIN 01 CMD2  $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$   
<sup>1</sup> This measurement is fully inclusive (includes “ $2e^+2e^-\gamma$ ” channel).

## $\Gamma(\pi^+\pi^-e^+e^- (\gamma))/\Gamma_{\text{total}}$ $\Gamma_{16}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.68±0.11 OUR FIT</b>				
<b>2.68±0.09±0.07</b>	1555 ±52	<sup>1</sup> AMBROSINO 09b	KLOE	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
4.3 +2.0-1.6 ±0.4 16 BERLOWSKI 08 WASA  $p d \rightarrow {}^3\text{He} \eta$   
4.3 ±1.3 ±0.4 16 BARGHOLTZ 07 CNTR See BERLOWSKI 08  
3.7 +2.5-1.8 ±0.3 4 AKHMETSHIN 01 CMD2  $e^+e^- \rightarrow \phi \rightarrow \eta\gamma$

<sup>1</sup> This AMBROSINO 09b value includes radiative events.

## $\Gamma(e^+e^-\mu^+\mu^-)/\Gamma_{\text{total}}$ $\Gamma_{17}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.6 × 10<sup>-4</sup></b>	90	BERLOWSKI 08	WASA	$p d \rightarrow {}^3\text{He} \eta$

## $\Gamma(2\mu^+2\mu^-)/\Gamma_{\text{total}}$ $\Gamma_{18}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.6 × 10<sup>-4</sup></b>	90	BERLOWSKI 08	WASA	$p d \rightarrow {}^3\text{He} \eta$

## $\Gamma(\mu^+\mu^-\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{19}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.6 × 10<sup>-4</sup></b>	90	BERLOWSKI 08	WASA	$p d \rightarrow {}^3\text{He} \eta$

## $\Gamma(\pi^+e^-\nu_e + \text{c.c.})/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_{20}/\Gamma_9$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.3 × 10<sup>-4</sup></b>	90	ABLIKIM 13g	BES3	$J/\psi \rightarrow \phi\eta$

## $\Gamma(\pi^+\pi^-2\gamma)/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_{21}/\Gamma_9$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 9 × 10<sup>-3</sup></b>		PRICE 67	HBC	

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<16 × 10<sup>-3</sup> 95 BALTAY 67b DBC

## $\Gamma(\pi^+\pi^-\pi^0\gamma)/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_{22}/\Gamma_9$

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.24 × 10<sup>-2</sup></b>	90	0	THALER 73	ASPK	

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<1.7 × 10<sup>-2</sup> 90 ARNOLD 68 HLBC  
<1.6 × 10<sup>-2</sup> 95 BALTAY 67b DBC  
<7.0 × 10<sup>-2</sup> 67 FLATTE 67 HBC  
<0.9 × 10<sup>-2</sup> 67 PRICE 67 HBC

## $\Gamma(\pi^0\mu^+\mu^-\gamma)/\Gamma_{\text{total}}$ $\Gamma_{23}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3 × 10<sup>-6</sup></b>	90	DZHELYADIN 81	SPEC	$\pi^- p \rightarrow \eta n$

### Forbidden modes

## $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$ $\Gamma_{24}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;9 × 10<sup>-5</sup></b>	90	NEFKENS 05a	CRYB	$p(720 \text{ MeV}/c) \pi^- \rightarrow n\eta$

Forbidden by angular momentum conservation.

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
Forbidden by $P$ and $CP$ invariance.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 0.13 \times 10^{-4}$	90	16M	AMBROSINO 05A	KLOE	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 3.9 \times 10^{-4}$	90	225M	ABLIKIM 11G	BES3	$e^+e^- \rightarrow J/\psi \rightarrow \eta\gamma$
$< 3.3 \times 10^{-4}$	90		AKHMETSHIN 99B	CMD2	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
$< 9 \times 10^{-4}$	90		AKHMETSHIN 97C	CMD2	See AKHMETSHIN 99B
$< 15 \times 10^{-4}$		0	THALER 73	ASPK	

$\Gamma(2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
Forbidden by $P$ and $CP$ invariance.					
VALUE	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
$<3.5 \times 10^{-4}$	90		BLIK 07	GAM4	$\pi^-p \rightarrow \eta n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<6.9 \times 10^{-4}$	90	225M	ABLIKIM 11G	BES3	$e^+e^- \rightarrow J/\psi \rightarrow \eta\gamma$
$<4.3 \times 10^{-4}$	90		AKHMETSHIN 99c	CMD2	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
$<6 \times 10^{-4}$	90		<sup>1</sup> ACHASOV 98	SND	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$
<sup>1</sup> ACHASOV 98 observes one event in a $\pm 3\sigma$ region around the $\eta$ mass, while a Monte Carlo calculation gives $10 \pm 5$ events. The limit here is the Poisson upper limit for one observed event and no background.					

$\Gamma(2\pi^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
Forbidden by C invariance.					
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
$< 5 \times 10^{-4}$	90	NEFKENS 05	CRYB 0		$p(720 \text{ MeV}/c) \pi^- \rightarrow n\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 17 \times 10^{-4}$	90	BLIK 07	GAM4		$\pi^- p \rightarrow \eta n$

$\Gamma(3\pi^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
Forbidden by C invariance.					
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
$< 6 \times 10^{-5}$	90	NEFKENS 05	CRYB 0		$p(720 \text{ MeV}/c) \pi^- \rightarrow n\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 24 \times 10^{-5}$	90	BLIK 07	GAM4		$\pi^- p \rightarrow \eta n$

$\Gamma(3\gamma)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
Forbidden by C invariance.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 16 \times 10^{-5}$	90	BLIK	07	GAM4	$\pi^- p \rightarrow \eta n$
$< 4 \times 10^{-5}$	90	NEFKENS	05A	CRYB	$p(720 \text{ MeV}/c) \pi^- \rightarrow n\eta$

$\Gamma(3\gamma)/\Gamma(2\gamma)$					$\Gamma_{29}/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	CHG		
$< 1.2 \times 10^{-3}$	95	ALDE 84	GAM2 0			

$\Gamma(3\gamma)/\Gamma(3\pi^0)$					$\Gamma_{29}/\Gamma_3$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$< 4.9 \times 10^{-5}$	90	ALOISIO 04	KLOE		$\phi \rightarrow \eta\gamma$	

$\Gamma(4\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
Forbidden by $P$ and $CP$ invariance.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 6.9 \times 10^{-7}$	90	PRAKHOV 00	CRYB	$\pi^- p \rightarrow n\eta$ , 720 MeV/c	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 200 \times 10^{-7}$	90	BLIK 07	GAM4	$\pi^- p \rightarrow \eta n$	

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$				$\Gamma_{31}/\Gamma$
C parity forbids this to occur as a single-photon process.				
VALUE	CL%	DOCUMENT ID	TECN	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.6 \times 10^{-4}$	90	MARTYNOV 76	HLBC	
$< 8.4 \times 10^{-4}$	90	BAZIN 68	DBC	
$< 70 \times 10^{-4}$		RITTENBERG 65	HBC	

$\Gamma(\pi^0 e^+ e^-)/\Gamma(\pi^+ \pi^- \pi^0)$					$\Gamma_{31}/\Gamma_9$	
C parity forbids this to occur as a single-photon process.						
VALUE	CL%	EVTS	DOCUMENT ID	TECN		
$< 1.9 \times 10^{-4}$	90		JANE 75	OSPK		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$< 42 \times 10^{-4}$	90		BAGLIN 67	HLBC		
$< 16 \times 10^{-4}$	90	0	BILLING 67	HLBC		
$< 77 \times 10^{-4}$		0	FOSTER 65B	HBC		
$< 110 \times 10^{-4}$			PRICE 65	HBC		

$\Gamma(\pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$				$\Gamma_{32}/\Gamma$
C parity forbids this to occur as a single-photon process.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 5 \times 10^{-6}$	90	DZHELYADIN 81	SPEC	$\pi^- p \rightarrow \eta n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 500 \times 10^{-6}$		WEH MANN 68	OSP K	

$[\Gamma(\mu^+ e^-) + \Gamma(\mu^- e^+)]/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$
Forbidden by lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6 \times 10^{-6}$	90	WHITE 96	SPEC	$p d \rightarrow \eta^3\text{He}$	

### $\eta$ C-NONCONSERVING DECAY PARAMETERS

**$\pi^+\pi^-\pi^0$  LEFT-RIGHT ASYMMETRY PARAMETER**  
Measurements with an error  $> 1.0 \times 10^{-2}$  have been omitted.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN
<b><math>0.09 \pm 0.11</math></b>	<b>OUR AVERAGE</b>		
$+0.09 \pm 0.10$	$+0.09$	1.34M	AMBROSINO 08D KLOE
$0.28 \pm 0.26$	165k	JANE 74	OSPK
$-0.05 \pm 0.22$	220k	LAYER 72	ASPK
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.5 \pm 0.5$	37k	<sup>1</sup> GORMLEY 68C	ASPK
<sup>1</sup> The GORMLEY 68C asymmetry is probably due to unmeasured ( $\mathbf{E} \times \mathbf{B}$ ) spark chamber effects. New experiments with ( $\mathbf{E} \times \mathbf{B}$ ) controls don't observe an asymmetry.			

**$\pi^+\pi^-\pi^0$  SEXTANT ASYMMETRY PARAMETER**  
Measurements with an error  $> 2.0 \times 10^{-2}$  have been omitted.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN
<b><math>0.12 \pm 0.10</math></b>	<b>OUR AVERAGE</b>		
$+0.08 \pm 0.10$	$+0.08$	1.34M	AMBROSINO 08D KLOE
$0.20 \pm 0.25$	165k	JANE 74	OSPK
$0.10 \pm 0.22$	220k	LAYER 72	ASPK
$0.5 \pm 0.5$	37k	GORMLEY 68C	WIRE

### $\pi^+\pi^-\pi^0$ QUADRANT ASYMMETRY PARAMETER

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN
<b><math>-0.09 \pm 0.09</math></b>	<b>OUR AVERAGE</b>		
$-0.05 \pm 0.10$	$+0.03$	1.34M	AMBROSINO 08D KLOE
$-0.30 \pm 0.25$	165k	JANE 74	OSPK
$-0.07 \pm 0.22$	220k	LAYER 72	ASPK

**$\pi^+\pi^-\gamma$  LEFT-RIGHT ASYMMETRY PARAMETER**  
Measurements with an error  $> 2.0 \times 10^{-2}$  have been omitted.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN
<b><math>0.9 \pm 0.4</math></b>	<b>OUR AVERAGE</b>		
$1.2 \pm 0.6$	35k	JANE 74B	OSPK
$0.5 \pm 0.6$	36k	THALER 72	ASPK
$1.22 \pm 1.56$	7257	GORMLEY 70	ASPK

### $\pi^+\pi^-\gamma$ PARAMETER $\beta$ ( $D$ -wave)

Sensitive to a  $D$ -wave contribution:  $dN/d\cos\theta = \sin^2\theta (1 + \beta \cos^2\theta)$ .

VALUE	EVTS	DOCUMENT ID	TECN
<b><math>-0.02 \pm 0.07</math></b>	<b>OUR AVERAGE</b>		
$0.11 \pm 0.11$	35k	JANE 74B	OSPK
$-0.060 \pm 0.065$	7250	GORMLEY 70	WIRE
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.12 \pm 0.06$		<sup>1</sup> THALER 72	ASPK

<sup>1</sup>The authors don't believe this indicates  $D$ -wave because the dependence of  $\beta$  on the  $\gamma$  energy is inconsistent with the theoretical prediction. A  $\cos^2\theta$  dependence can also come from  $P$ - and  $F$ -wave interference.

### $\eta$ CP-NONCONSERVING DECAY PARAMETER

#### $\pi^+\pi^-e^+e^-$ DECAY-PLANE ASYMMETRY PARAMETER $A_\phi$

In the  $\eta$  rest frame, the total momentum of the  $e^+e^-$  pair is equal and opposite to that of the  $\pi^+\pi^-$  pair. Let  $\hat{z}$  be the unit vector along the momentum of the  $e^+e^-$  pair; let  $\hat{n}_{ee}$  and  $\hat{n}_{\pi\pi}$  be the unit vectors normal to the  $e^+e^-$  and  $\pi^+\pi^-$  planes; and let  $\phi$  be the angle between the two normals. Then

$$\sin\phi \cos\phi = [(\hat{n}_{ee} \times \hat{n}_{\pi\pi}) \cdot \hat{z}] (\hat{n}_{ee} \cdot \hat{n}_{\pi\pi}),$$

and

$$A_\phi \equiv \frac{N_{\sin\phi \cos\phi > 0} - N_{\sin\phi \cos\phi < 0}}{N_{\sin\phi \cos\phi > 0} + N_{\sin\phi \cos\phi < 0}}.$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.6 \pm 2.5 \pm 1.8</math></b>	$1555 \pm 52$	AMBROSINO 09B	KLOE	$e^+e^- \rightarrow \phi \rightarrow \eta\gamma$

### ENERGY DEPENDENCE OF $\eta \rightarrow 3\pi$ DALITZ PLOTS

#### PARAMETERS FOR $\eta \rightarrow \pi^+\pi^-\pi^0$

See the "Note on  $\eta$  Decay Parameters" in our 1994 edition, Phys. Rev. **D50**, 1 August 1994, Part I, p. 1454. The following experiments fit to one or more of the coefficients  $a, b, c, d$ , or  $e$  for  $|\text{matrix element}|^2 = 1 + ay + by^2 + cx + dx^2 + exy$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
	79k	ABLIKIM 15G	BES3	$e^+e^- \rightarrow J/\psi \rightarrow \eta\gamma$



$f_0(500)$  or  $\sigma$   
was  $f_0(600)$

$$J^G(J^{PC}) = 0^+(0^{++})$$

## NOTE ON SCALAR MESONS BELOW 2 GEV

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**I. Introduction:** In contrast to the vector and tensor mesons, the identification of the scalar mesons is a long-standing puzzle. Scalar resonances are difficult to resolve because some of them have large decay widths which cause a strong overlap between resonances and background. In addition, several decay channels sometimes open up within a short mass interval (*e.g.* at the  $K\bar{K}$  and  $\eta\eta$  thresholds), producing cusps in the line shapes of the near-by resonances. Furthermore, one expects non- $q\bar{q}$  scalar objects, such as glueballs and multi-quark states in the mass range below 2 GeV (for reviews see, *e.g.*, Refs. [1–5] and the mini-review on *non- $q\bar{q}$  states* in this Review of Particle Physics (RPP)).

Light scalars are produced, for example, in  $\pi N$  scattering on polarized/unpolarized targets,  $p\bar{p}$  annihilation, central hadronic production,  $J/\Psi$ ,  $B^-$ ,  $D^-$  and  $K$ -meson decays,  $\gamma\gamma$  formation, and  $\phi$  radiative decays. Especially for the lightest scalar mesons simple parameterizations fail and more advanced theory tools are necessary to extract the resonance parameters from data. In the analyses available in the literature fundamental properties of the amplitudes such as unitarity, analyticity, Lorentz invariance, chiral and flavor symmetry are implemented at different levels of rigor. Especially, chiral symmetry implies the appearance of zeros close to the threshold in elastic  $S$ -wave scattering amplitudes involving soft pions [6,7], which may be shifted or removed in associated production processes [8]. The methods employed are the  $K$ -matrix formalism, the  $N/D$ -method, the Dalitz–Tuan ansatz, unitarized quark models with coupled channels, effective chiral field theories and the linear sigma model, *etc.* Dynamics near the lowest two-body thresholds in some analyses are described by crossed channel ( $t$ ,  $u$ ) meson exchange or with an effective range parameterization instead of, or in addition to, resonant features in the  $s$ -channel. Dispersion theoretical approaches are applied to pin down the location of resonance poles for the low-lying states [9–12].

The mass and width of a resonance are found from the position of the nearest pole in the process amplitude ( $T$ -matrix or  $S$ -matrix) at an unphysical sheet of the complex energy plane, traditionally labeled as

$$\sqrt{s_{\text{Pole}}} = M - i\Gamma/2.$$

It is important to note that the pole of a Breit-Wigner parameterization agrees with this pole position only for narrow

and well-separated resonances, far away from the opening of decay channels. For a detailed discussion of this issue we refer to the review on *Resonances* in this RPP.

In this note, we discuss the light scalars below 2 GeV organized in the listings under the entries ( $I = 1/2$ )  $K_0^*(800)$  (or  $\kappa$ , currently omitted from the summary table),  $K_0^*(1430)$ , ( $I = 1$ )  $a_0(980)$ ,  $a_0(1450)$ , and ( $I = 0$ )  $f_0(500)$  (or  $\sigma$ ),  $f_0(980)$ ,  $f_0(1370)$ ,  $f_0(1500)$ , and  $f_0(1710)$ . This list is minimal and does not necessarily exhaust the list of actual resonances. The ( $I = 2$ )  $\pi\pi$  and ( $I = 3/2$ )  $K\pi$  phase shifts do not exhibit any resonant behavior.

**II. The  $I = 1/2$  States:** The  $K_0^*(1430)$  [13] is perhaps the least controversial of the light scalar mesons. The  $K\pi$   $S$ -wave scattering has two possible isospin channels,  $I = 1/2$  and  $I = 3/2$ . The  $I = 3/2$  wave is elastic and repulsive up to 1.7 GeV [14] and contains no known resonances. The  $I = 1/2$   $K\pi$  phase shift, measured from about 100 MeV above threshold in  $Kp$  production, rises smoothly, passes  $90^\circ$  at 1350 MeV, and continues to rise to about  $170^\circ$  at 1600 MeV. The first important inelastic threshold is  $K\eta'(958)$ . In the inelastic region the continuation of the amplitude is uncertain since the partial-wave decomposition has several solutions. The data are extrapolated towards the  $K\pi$  threshold using effective range type formulas [13,15] or chiral perturbation predictions [16,17]. From analyses using unitarized amplitudes there is agreement on the presence of a resonance pole around 1410 MeV having a width of about 300 MeV. With reduced model dependence, Ref. [18] finds a larger width of 500 MeV.

Similar to the situation for the  $f_0(500)$ , discussed in the next section, the presence and properties of the light  $K_0^*(800)$  (or  $\kappa$ ) meson in the 700–900 MeV region are difficult to establish since it appears to have a very large width ( $\Gamma \approx 500$  MeV) and resides close to the  $K\pi$  threshold. Hadronic  $D$ - and  $B$ -meson decays provide additional data points in the vicinity of the  $K\pi$  threshold and are discussed in detail in the *Review on Multibody Charm Analyses* in this RPP. Precision information from semileptonic  $D$  decays avoiding theoretically ambiguous three-body final state interactions is not available. BES II [19] (re-analyzed in [20]) finds a  $K_0^*(800)$ -like structure in  $J/\psi$  decays to  $\bar{K}^{*0}(892)K^+\pi^-$  where  $K_0^*(800)$  recoils against the  $K^*(892)$ . Also clean with respect to final state interaction is the decay  $\tau^- \rightarrow K_S^0\pi^-\nu_\tau$  studied by Belle [21], with  $K_0^*(800)$  parameters fixed to those of Ref. [19].

Some authors find a  $K_0^*(800)$  pole in their phenomenological analysis (see, *e.g.*, [22–33]), while others do not need to include it in their fits (see, *e.g.*, [17,34–37]). Similarly to the case of the  $f_0(500)$  discussed below, all works including constraints from chiral symmetry at low energies naturally seem to find a light  $K_0^*(800)$  below 800 MeV, see, *e.g.*, [38–42]. In these works the  $K_0^*(800)$ ,  $f_0(500)$ ,  $f_0(980)$  and  $a_0(980)$  appear to form a nonet [39,40]. Additional evidence for this assignment is presented in Ref. [12], where the couplings of the nine states to  $q\bar{q}$  sources were compared. The same low-lying scalar nonet was

# Meson Particle Listings

## $f_0(500)$

also found earlier in the unitarized quark model of Ref. [41]. The analysis of Ref. [43] is based on the Roy-Steiner equations, which include analyticity and crossing symmetry. It establishes the existence of a light  $K_0^*(800)$  pole in the  $K\pi \rightarrow K\pi$  amplitude on the second sheet. In Ref. [44] a first lattice study for the  $K\pi$  S-wave system is presented, however, with a pion mass of 400 MeV it can not be compared to data yet.

**III. The  $I = 1$  States:** Two isovector scalar states are known below 2 GeV, the  $a_0(980)$  and the  $a_0(1450)$ . Independent of any model, the  $K\bar{K}$  component in the  $a_0(980)$  wave function must be large: it lies just below the opening of the  $K\bar{K}$  channel to which it strongly couples [15,45]. This generates an important cusp-like behavior in the resonant amplitude. Hence, its mass and width parameters are strongly distorted. To reveal its true coupling constants, a coupled-channel model with energy-dependent widths and mass shift contributions is necessary. All listed  $a_0(980)$  measurements agree on a mass position value near 980 MeV, but the width takes values between 50 and 100 MeV, mostly due to the different models. For example, the analysis of the  $p\bar{p}$ -annihilation data [15] using a unitary  $K$ -matrix description finds a width as determined from the  $T$ -matrix pole of  $92 \pm 8$  MeV, while the observed width of the peak in the  $\pi\eta$  mass spectrum is about 45 MeV.

The relative coupling  $K\bar{K}/\pi\eta$  is determined indirectly from  $f_1(1285)$  [46–48] or  $\eta(1410)$  decays [49–51], from the line shape observed in the  $\pi\eta$  decay mode [52–55], or from the coupled-channel analysis of the  $\pi\pi\eta$  and  $K\bar{K}\pi$  final states of  $p\bar{p}$  annihilation at rest [15].

The  $a_0(1450)$  is seen in  $p\bar{p}$  annihilation experiments with stopped and higher momenta antiprotons, with a mass of about 1450 MeV or close to the  $a_2(1320)$  meson which is typically a dominant feature. A contribution from  $a_0(1450)$  is also found in the analysis of the  $D^\pm \rightarrow K^+ K^- \pi^\pm$  decay [56]. The broad structure at about 1300 MeV observed in  $\pi N \rightarrow K\bar{K}N$  reactions [57] needs still further confirmation in its existence and isospin assignment.

**IV. The  $I = 0$  States:** The  $I = 0$ ,  $J^{PC} = 0^{++}$  sector is the most complex one, both experimentally and theoretically. The data have been obtained from the  $\pi\pi$ ,  $K\bar{K}$ ,  $\eta\eta$ ,  $4\pi$ , and  $\eta\eta'(958)$  systems produced in  $S$ -wave. Analyses based on several different production processes conclude that probably four poles are needed in the mass range from  $\pi\pi$  threshold to about 1600 MeV. The claimed isoscalar resonances are found under separate entries  $f_0(500)$  (or  $\sigma$ ),  $f_0(980)$ ,  $f_0(1370)$ , and  $f_0(1500)$ .

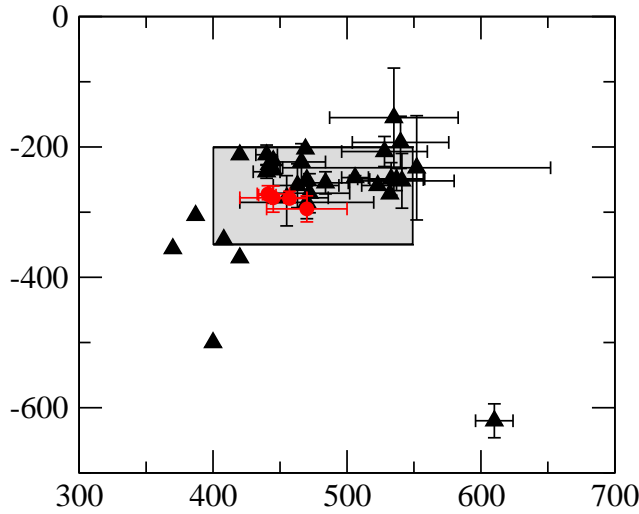
For discussions of the  $\pi\pi$   $S$  wave below the  $K\bar{K}$  threshold and on the long history of the  $f_0(500)$ , which was suggested in linear sigma models more than 50 years ago, see our reviews in previous editions and the recent review [5].

Information on the  $\pi\pi$   $S$ -wave phase shift  $\delta_f^I = \delta_0^0$  was already extracted many years ago from  $\pi N$  scattering [58–60],

and near threshold from the  $K_{e4}$ -decay [61]. The kaon decays were later revisited leading to consistent data, however, with very much improved statistics [62,63]. The reported  $\pi\pi \rightarrow K\bar{K}$  cross sections [64–67] have large uncertainties. The  $\pi N$  data have been analyzed in combination with high-statistics data (see entries labeled as RVUE for re-analyses of the data). The  $2\pi^0$  invariant mass spectra of the  $p\bar{p}$  annihilation at rest [68–70] and the central collision [71] do not show a distinct resonance structure below 900 MeV, but these data are consistently described with the standard solution for  $\pi N$  data [59,72], which allows for the existence of the broad  $f_0(500)$ . An enhancement is observed in the  $\pi^+\pi^-$  invariant mass near threshold in the decays  $D^+ \rightarrow \pi^+\pi^-\pi^+$  [73–101] and  $J/\psi \rightarrow \omega\pi^+\pi^-$  [76,98], and in  $\psi(2S) \rightarrow J/\psi\pi^+\pi^-$  with very limited phase space [78,79].

The precise  $f_0(500)$  (or  $\sigma$ ) pole is difficult to establish because of its large width, and because it can certainly not be modeled by a naive Breit-Wigner resonance. For the same reason a splitting in background and resonance contributions is not possible in a model-independent way. The  $\pi\pi$  scattering amplitude shows an unusual energy dependence due to the presence of a zero in the unphysical regime close to the threshold [6–7], required by chiral symmetry, and possibly due to crossed channel exchanges, the  $f_0(1370)$ , and other dynamical features. However, most of the analyses listed under  $f_0(500)$  agree on a pole position near  $(500 - i250)$  MeV. In particular, analyses of  $\pi\pi$  data that include unitarity,  $\pi\pi$  threshold behavior, strongly constrained by the  $K_{e4}$  data, and the chiral symmetry constraints from Adler zeroes and/or scattering lengths find a light  $f_0(500)$ , see, *e.g.*, [80,81].

Precise pole positions with an uncertainty of less than 20 MeV (see our table for the  $T$ -matrix pole) were extracted by use of Roy equations, which are twice subtracted dispersion relations derived from crossing symmetry and analyticity. In Ref. [10] the subtraction constants were fixed to the  $S$ -wave scattering lengths  $a_0^0$  and  $a_2^0$  derived from matching Roy equations and two-loop chiral perturbation theory [9]. The only additional relevant input to fix the  $f_0(500)$  pole turned out to be the  $\pi\pi$ -wave phase shifts at 800 MeV. The analysis was improved further in Ref. [12]. Alternatively, in Ref. [11] only data were used as input inside Roy equations. In that reference also once-subtracted Roy-like equations, called GKPY equations, were used, since the extrapolation into the complex plane based on the twice subtracted equations leads to larger uncertainties mainly due to the limited experimental information on the isospin-2  $\pi\pi$  scattering length. All these extractions find consistent results. Using analyticity and unitarity only to describe data from  $K_{2\pi}$  and  $K_{e4}$  decays, Ref. [82] finds consistent values for the pole position and the scattering length  $a_0^0$ . The importance of the  $\pi\pi$  scattering data for fixing the  $f_0(500)$  pole is nicely illustrated by comparing analyses of  $p\bar{p} \rightarrow 3\pi^0$  omitting [68,83] or including [69,84] information on  $\pi\pi$  scattering: while the former analyses find an extremely broad structure above 1 GeV, the latter find  $f_0(500)$  masses of the order of 400 MeV.



**Figure 1:** Location of the  $f_0(500)$  (or  $\sigma$ ) poles in the complex energy plane. Circles denote the recent analyses based on Roy(-like) dispersion relations [9–12], while all other analyses are denoted by triangles. The corresponding references are given in the listing.

As a result of the sensitivity of the extracted  $f_0(500)$  pole position on the high accuracy low energy  $\pi\pi$  scattering data [62,63], the currently quoted range of pole positions for the  $f_0(500)$ , namely

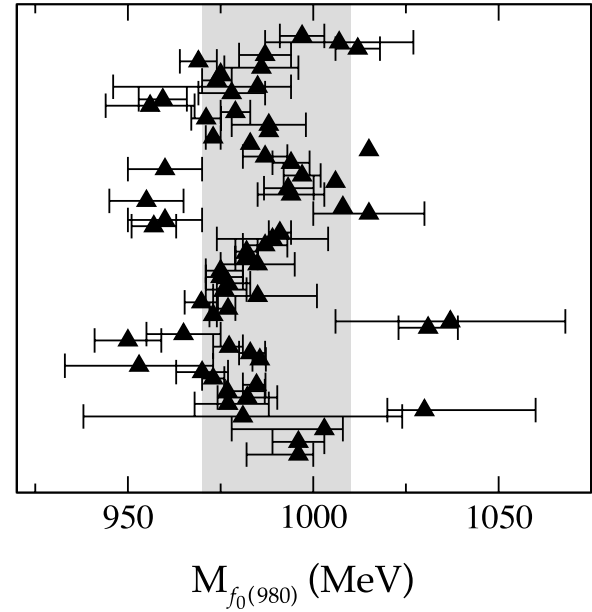
$$\sqrt{s_{\text{Pole}}} = (400 - 550) - i(200 - 350) \text{ MeV},$$

in the listing was fixed including only those analyses consistent with these data, Refs. [26,29,39,41,42,54,69,78–82,85–101] as well as the advanced dispersion analyses [9–12]. The pole positions from those references are compared to the range of pole positions quoted above in Fig. 1. Note that this range is labeled as 'our estimate' — it is not an average over the quoted analyses but is chosen to include the bulk of the analyses consistent with the mentioned criteria. An averaging procedure is not justified, since the analyses use overlapping or identical data sets.

One might also take the more radical point of view and just average the most advanced dispersive analyses, Refs. [9–12], shown as solid dots in Fig. 1, for they provide a determination of the pole positions with minimal bias. This procedure leads to the much more restricted range of  $f_0(500)$  parameters

$$\sqrt{s_{\text{Pole}}} = (446 \pm 6) - i(276 \pm 5) \text{ MeV}.$$

Due to the large strong width of the  $f_0(500)$  an extraction of its two-photon width directly from data is not possible. Thus, the values for  $\Gamma(\gamma\gamma)$  quoted in the literature as well as the listing are based on the expression in the narrow width approximation [102]  $\Gamma(\gamma\gamma) \simeq \alpha^2 |g_\gamma|^2 / (4\text{Re}(\sqrt{s_{\text{Pole}}}))$  where  $g_\gamma$  is derived from the residue at the  $f_0(500)$  pole to two photons and  $\alpha$  denotes the electromagnetic fine structure constant.



**Figure 2:** Values of the  $f_0(980)$  masses as they appear in the listing compared to the currently quoted mass estimate. The newest references appear at the bottom, the oldest on the top. The corresponding references are given in the listing.

The explicit form of the expression may vary between different authors due to different definitions of the coupling constant, however, the expression given for  $\Gamma(\gamma\gamma)$  is free of ambiguities. According to Refs. [103,104], the data for  $f_0(500) \rightarrow \gamma\gamma$  are consistent with what is expected for a two-step process of  $\gamma\gamma \rightarrow \pi^+\pi^-$  via pion exchange in the  $t$ - and  $u$ -channel, followed by a final state interaction  $\pi^+\pi^- \rightarrow \pi^0\pi^0$ . The same conclusion is drawn in Ref. [105] where the bulk part of the  $f_0(500) \rightarrow \gamma\gamma$  decay width is dominated by re-scattering. Therefore, it might be difficult to learn anything new about the nature of the  $f_0(500)$  from its  $\gamma\gamma$  coupling. For the most recent work on  $\gamma\gamma \rightarrow \pi\pi$ , see [106–108]. There are theoretical indications (*e.g.*, [109–112]) that the  $f_0(500)$  pole behaves differently from a  $q\bar{q}$ -state – see next section and the mini-review on *non  $q\bar{q}$ -states* in this RPP for details.

The  **$f_0(980)$**  overlaps strongly with the background represented mainly by the  $f_0(500)$  and the  $f_0(1370)$ . This can lead to a dip in the  $\pi\pi$  spectrum at the  $K\bar{K}$  threshold. It changes from a dip into a peak structure in the  $\pi^0\pi^0$  invariant mass spectrum of the reaction  $\pi^-p \rightarrow \pi^0\pi^0n$  [113], with increasing four-momentum transfer to the  $\pi^0\pi^0$  system, which means increasing the  $a_1$ -exchange contribution in the amplitude, while the  $\pi$ -exchange decreases. The  $f_0(500)$  and the  $f_0(980)$  are also observed in data for radiative decays ( $\phi \rightarrow f_0\gamma$ ) from SND [114,115], CMD2 [116], and KLOE [117,118]. A dispersive analysis was used to simultaneously pin down the pole parameters of both the  $f_0(500)$  and the  $f_0(980)$  [11]; the uncertainty in the pole position quoted for the latter state is

# Meson Particle Listings

## $f_0(500)$

of the order of 10 MeV, only (see the lowest point in Fig. 2). Compared to the 2010 issue of the *Review of Particle Physics*, in this issue we extended the allowed range of the  $f_0(980)$  masses to include the mass value derived in Ref. [11]. We now quote for the mass

$$M_{f_0(980)} = 990 \pm 20 \text{ MeV}.$$

As in case of the  $f_0(500)$  (or  $\sigma$ ), this range is not an average, but is labeled as 'our estimate'. A comparison of the mass values in the listing and the allocated range is shown in Fig. 2.

Analyses of  $\gamma\gamma \rightarrow \pi\pi$  data [119–121] underline the importance of the  $K\bar{K}$  coupling of  $f_0(980)$ , while the resulting two-photon width of the  $f_0(980)$  cannot be determined precisely [122]. The prominent appearance of the  $f_0(980)$  in the semileptonic  $D_s$  decays and decays of  $B$  and  $B_s$ -mesons implies a dominant ( $\bar{s}s$ ) component: those decays occur via weak transitions that alternatively result in  $\phi(1020)$  production. Ratios of decay rates of  $B$  and/or  $B_s$  mesons into  $J/\psi$  plus  $f_0(980)$  or  $f_0(500)$  were proposed to allow for an extraction of the flavor mixing angle and to probe the tetraquark nature of those mesons within a certain model [123,124]. The phenomenological fits of the LHCb collaboration using the isobar model do neither allow for a contribution of the  $f_0(980)$  in the  $B \rightarrow J/\psi\pi\pi$  [125] nor for an  $f_0(500)$  in  $B_s \rightarrow J/\psi\pi\pi$  decays [126]. From the former analysis the authors conclude that their data is incompatible with a model where  $f_0(500)$  and  $f_0(980)$  are formed from two quarks and two antiquarks (tetraquarks) at the eight standard deviation level. In addition, they extract an upper limit for the mixing angle of  $17^\circ$  at 90% C.L. between the  $f_0(980)$  and the  $f_0(500)$  that would correspond to a substantial ( $\bar{s}s$ ) content in  $f_0(980)$  [125]. However, in a dispersive analysis of the same data that allows for a model-independent inclusion of the hadronic final state interactions in Ref. [127] a substantial  $f_0(980)$  contribution is also found in the  $B$ -decays putting into question the conclusions of Ref. [125].

**The  $f_0$ 's above 1 GeV.** A meson resonance that is very well studied experimentally, is the  **$f_0(1500)$**  seen by the Crystal Barrel experiment in five decay modes:  $\pi\pi$ ,  $K\bar{K}$ ,  $\eta\eta$ ,  $\eta\eta'(958)$ , and  $4\pi$  [15,69,70]. Due to its interference with the  $f_0(1370)$  (and  **$f_0(1710)$** ), the peak attributed to the  $f_0(1500)$  can appear shifted in invariant mass spectra. Therefore, the application of simple Breit-Wigner forms arrives at slightly different resonance masses for  $f_0(1500)$ . Analyses of central-production data of the likewise five decay modes Refs. [128,129] agree on the description of the  $S$ -wave with the one above. The  $p\bar{p}$ ,  $p\bar{n}/n\bar{p}$  measurements [70,130–132] show a single enhancement at 1400 MeV in the invariant  $4\pi$  mass spectra, which is resolved into  $f_0(1370)$  and  $f_0(1500)$  [133,134]. The data on  $4\pi$  from central production [135] require both resonances, too, but disagree on the relative content of  $\rho\rho$  and  $f_0(500)f_0(500)$  in  $4\pi$ . All investigations agree that the  $4\pi$  decay mode represents about half of the  $f_0(1500)$  decay width and is dominant for  $f_0(1370)$ .

The determination of the  $\pi\pi$  coupling of  **$f_0(1370)$**  is aggravated by the strong overlap with the broad  $f_0(500)$  and  $f_0(1500)$ . Since it does not show up prominently in the  $2\pi$  spectra, its mass and width are difficult to determine. Multi-channel analyses of hadronically produced two- and three-body final states agree on a mass between 1300 MeV and 1400 MeV and a narrow  $f_0(1500)$ , but arrive at a somewhat smaller width for  $f_0(1370)$ .

**V. Interpretation of the scalars below 1 GeV:** In the literature, many suggestions are discussed, such as conventional  $q\bar{q}$  mesons,  $q\bar{q}q\bar{q}$  or meson-meson bound states. In addition, one expects a scalar glueball in this mass range. In reality, there can be superpositions of these components, and one often depends on models to determine the dominant one. Although we have seen progress in recent years, this question remains open. Here, we mention some of the present conclusions.

The  $f_0(980)$  and  $a_0(980)$  are often interpreted as multiquark states [136–140] or  $K\bar{K}$  bound states [141]. The insight into their internal structure using two-photon widths [115,142–148] is not conclusive. The  $f_0(980)$  appears as a peak structure in  $J/\psi \rightarrow \phi\pi^+\pi^-$  and in  $D_s$  decays without  $f_0(500)$  background, while being nearly invisible in  $J/\psi \rightarrow \omega\pi^+\pi^-$ . Based on that observation it is suggested that  $f_0(980)$  has a large  $s\bar{s}$  component, which according to Ref. [149] is surrounded by a virtual  $K\bar{K}$  cloud (see also Ref. [150]). Data on radiative decays ( $\phi \rightarrow f_0\gamma$  and  $\phi \rightarrow a_0\gamma$ ) from SND, CMD2, and KLOE (see above) are consistent with a prominent role of kaon loops. This observation is interpreted as evidence for a compact four-quark [151] or a molecular [152,153] nature of these states. Details of this controversy are given in the comments [154,155]; see also Ref. [156]. It remains quite possible that the states  $f_0(980)$  and  $a_0(980)$ , together with the  $f_0(500)$  and the  $K_0^*(800)$ , form a new low-mass state nonet of predominantly four-quark states, where at larger distances the quarks recombine into a pair of pseudoscalar mesons creating a meson cloud (see, *e.g.*, Ref. [157]). Different QCD sum rule studies [158–162] do not agree on a tetraquark configuration for the same particle group.

Models that start directly from chiral Lagrangians, either in non-linear [42,25,80,152] or in linear [163–169] realization, predict the existence of the  $f_0(500)$  meson near 500 MeV. Here the  $f_0(500)$ ,  $a_0(980)$ ,  $f_0(980)$ , and  $K_0^*(800)$  (in some models the  $K_0^*(1430)$ ) would form a nonet (not necessarily  $q\bar{q}$ ). In the linear sigma models the lightest pseudoscalars appear as their chiral partners. In these models the light  $f_0(500)$  is often referred to as the "Higgs boson of strong interactions", since here the  $f_0(500)$  plays a role similar to the Higgs particle in electro-weak symmetry breaking: within the linear sigma models it is important for the mechanism of chiral symmetry breaking, which generates most of the proton mass, and what is referred to as the constituent quark mass.

In the non-linear approaches of [25,80] the above resonances together with the low lying vector states are generated starting from chiral perturbation theory predictions near the first open



channel, and then by extending the predictions to the resonance regions using unitarity and analyticity.

Ref. [163] uses a framework with explicit resonances that are unitarized and coupled to the light pseudoscalars in a chirally invariant way. Evidence for a non- $\bar{q}q$  nature of the lightest scalar resonances is derived from their mixing scheme. In Ref. [164] the scheme is extended and applied to the decay  $\eta' \rightarrow \eta\pi\pi$ , which lead to the same conclusions. To identify the nature of the resonances generated from scattering equations, in Ref. [170] the large  $N_c$  behavior of the poles was studied, with the conclusion that, while the light vector states behave consistent with what is predicted for  $\bar{q}q$  states, the light scalars behave very differently. This finding provides strong support for a non- $\bar{q}q$  nature of the light scalar resonances. Note, the more refined study of Ref. [109] found, in case of the  $f_0(500)$ , in addition to a dominant non- $\bar{q}q$  nature, indications for a subdominant  $\bar{q}q$  component located around 1 GeV. Additional support for the non- $\bar{q}q$  nature of the  $f_0(500)$  is given in Ref. [171], where the connection between the pole of resonances and their Regge trajectories is analyzed.

A model-independent method to identify hadronic molecules goes back to a proposal by Weinberg [172], shown to be equivalent to the pole counting arguments of [173–175] in Ref. [176]. The formalism allows one to extract the amount of molecular component in the wave function from the effective coupling constant of a physical state to a nearby continuum channel. It can be applied to near threshold states only and provided strong evidence that the  $f_0(980)$  is a  $\bar{K}K$  molecule, while the situation turned out to be less clear for the  $a_0(980)$  (see also Refs. [148,146]). Further insights into  $a_0(980)$  and  $f_0(980)$  are expected from their mixing [177]. The corresponding signal predicted in Refs. [178,179] was recently observed at BES III [180]. It turned out that in order to get a quantitative understanding of those data in addition to the mixing mechanism itself, some detailed understanding of the production mechanism seems necessary [181].

In the unitarized quark model with coupled  $q\bar{q}$  and meson-meson channels, the light scalars can be understood as additional manifestations of bare  $q\bar{q}$  confinement states, strongly mass shifted from the 1.3 - 1.5 GeV region and very distorted due to the strong  $^3P_0$  coupling to  $S$ -wave two-meson decay channels [182] vanbeveren01b. Thus, in these models the light scalar nonet comprising the  $f_0(500)$ ,  $f_0(980)$ ,  $K_0^*(800)$ , and  $a_0(980)$ , as well as the nonet consisting of the  $f_0(1370)$ ,  $f_0(1500)$  (or  $f_0(1710)$ ),  $K_0^*(1430)$ , and  $a_0(1450)$ , respectively, are two manifestations of the same bare input states (see also Ref. [184]).

Other models with different groupings of the observed resonances exist and may, *e.g.*, be found in earlier versions of this review.

**VI. Interpretation of the  $f_0$ 's above 1 GeV:** The  $f_0(1370)$  and  $f_0(1500)$  decay mostly into pions ( $2\pi$  and  $4\pi$ ) while the  $f_0(1710)$  decays mainly into the  $K\bar{K}$  final states. The  $K\bar{K}$  decay branching ratio of the  $f_0(1500)$  is small [128,185].

If one uses the naive quark model, it is natural to assume that the  $f_0(1370)$ ,  $a_0(1450)$ , and the  $K_0^*(1430)$  are in the same SU(3) flavor nonet, being the  $(u\bar{u} + d\bar{d})$ ,  $u\bar{d}$  and  $u\bar{s}$  states, probably mixing with the light scalars [186], while the  $f_0(1710)$  is the  $s\bar{s}$  state. Indeed, the production of  $f_0(1710)$  (and  $f_2'(1525)$ ) is observed in  $p\bar{p}$  annihilation [187] but the rate is suppressed compared to  $f_0(1500)$  (respectively,  $f_2(1270)$ ), as would be expected from the OZI rule for  $s\bar{s}$  states. The  $f_0(1500)$  would also qualify as a  $(u\bar{u} + d\bar{d})$  state, although it is very narrow compared to the other states and too light to be the first radial excitation.

However, in  $\gamma\gamma$  collisions leading to  $K_S^0 K_S^0$  [188] a spin-0 signal is observed at the  $f_0(1710)$  mass (together with a dominant spin-2 component), while the  $f_0(1500)$  is not observed in  $\gamma\gamma \rightarrow K\bar{K}$  nor  $\pi^+\pi^-$  [189]. In  $\gamma\gamma$  collisions leading to  $\pi^0\pi^0$  Ref. [190] reports the observation of a scalar around 1470 MeV albeit with large uncertainties on the mass and  $\gamma\gamma$  couplings. This state could be the  $f_0(1370)$  or the  $f_0(1500)$ . The upper limit from  $\pi^+\pi^-$  [189] excludes a large  $n\bar{n}$  (here  $n$  stands for the two lightest quarks) content for the  $f_0(1500)$  and hence points to a mainly  $s\bar{s}$  state [191]. This appears to contradict the small  $K\bar{K}$  decay branching ratio of the  $f_0(1500)$  and makes a  $q\bar{q}$  assignment difficult for this state. Hence the  $f_0(1500)$  could be mainly glue due the absence of a  $2\gamma$ -coupling, while the  $f_0(1710)$  coupling to  $2\gamma$  would be compatible with an  $s\bar{s}$  state. This is in accord with the recent high-statistics Belle data in  $\gamma\gamma \rightarrow K_S^0 K_S^0$  [192] in which the  $f_0(1500)$  is absent, while a prominent peak at 1710 MeV is observed with quantum numbers  $0^{++}$ , compatible with the formation of an  $s\bar{s}$  state. However, the  $2\gamma$ -couplings are sensitive to glue mixing with  $q\bar{q}$  [193].

Note that an isovector scalar, possibly the  $a_0(1450)$  (albeit at a lower mass of 1317 MeV) is observed in  $\gamma\gamma$  collisions leading to  $\eta\pi^0$  [194]. The state interferes destructively with the non-resonant background, but its  $\gamma\gamma$  coupling is comparable to that of the  $a_2(1320)$ , in accord with simple predictions (see, *e.g.*, Ref. [191]).

The small width of  $f_0(1500)$ , and its enhanced production at low transverse momentum transfer in central collisions [195–197] also favor  $f_0(1500)$  to be non- $q\bar{q}$ . In the mixing scheme of Ref. [193], which uses central production data from WA102 and the recent hadronic  $J/\psi$  decay data from BES [198,199], glue is shared between  $f_0(1370)$ ,  $f_0(1500)$  and  $f_0(1710)$ . The  $f_0(1370)$  is mainly  $n\bar{n}$ , the  $f_0(1500)$  mainly glue and the  $f_0(1710)$  dominantly  $s\bar{s}$ . This agrees with previous analyses [200,201]. In Ref. [202]  $f_0(1710)$  qualifies as a glueball candidate based on an analysis of decay data in an extended linear sigma model with a dilaton field.

However, alternative schemes have been proposed (*e.g.*, in [203–204]; for a review see, *e.g.*, Ref. [1]) . In particular, for a scalar glueball, the two-gluon coupling to  $n\bar{n}$  appears to be suppressed by chiral symmetry [205] and therefore the  $K\bar{K}$  decay could be enhanced. This mechanism would imply that the  $f_0(1710)$  can possibly be interpreted as an unmixed

# Meson Particle Listings

## $f_0(500)$

glueball [206]. In Ref. [207], a large  $K^+K^-$  scalar signal reported by Belle in  $B$  decays into  $KK\bar{K}$  [208], compatible with the  $f_0(1500)$ , is explained as due to constructive interference with a broad glueball background. However, the Belle data are inconsistent with the BaBar measurements which show instead a broad scalar at this mass for  $B$  decays into both  $K^\pm K^\pm K^\mp$  [209] and  $K^+K^-\pi^0$  [210].

Whether the  $f_0(1500)$  is observed in 'gluon rich' radiative  $J/\psi$  decays is debatable [211] because of the limited amount of data - more data for this and the  $\gamma\gamma$  mode are needed.

In Ref. [212], further refined in Ref. [213],  $f_0(1370)$  and  $f_0(1710)$  (together with  $f_2(1270)$  and  $f_2'(1525)$ ) were interpreted as bound systems of two vector mesons. This picture could be tested in radiative  $J/\psi$  decays [214] as well as radiative decays of the states themselves [215]. The vector-vector component of the  $f_0(1710)$  might also be the origin of the enhancement seen in  $J/\psi \rightarrow \gamma\phi\omega$  near threshold [216] observed at BES [217].

### References

1. C. Amsler and N.A. Tornqvist, Phys. Reports **389**, 61 (2004).
2. D.V. Bugg, Phys. Reports **397**, 257 (2004).
3. F.E. Close and N.A. Tornqvist, J. Phys. **G28**, R249 (2002).
4. E. Klempt and A. Zaitsev, Phys. Reports **454**, 1 (2007).
5. J.R. Pelaez, [arXiv:1510.00653 \[hep-ph\]](#).
6. J.L. Adler, Phys. Rev. **137**, B1022 (1965).
7. J.L. Adler, Phys. Rev. **139**, B1638 (1965).
8. J.A. Oller, Phys. Rev. **D71**, 054030 (2005).
9. G. Colangelo, J. Gasser, and H. Leutwyler, Nucl. Phys. **B603**, 125 (2001).
10. I. Caprini, G. Colangelo, and H. Leutwyler, Phys. Rev. Lett. **96**, 132001 (2006).
11. R. Garcia-Martin, *et al.*, Phys. Rev. Lett. **107**, 072001 (2011).
12. B. Moussallam, Eur. Phys. J. **C71**, 1814 (2011).
13. D. Aston, *et al.*, Nucl. Phys. **B296**, 493 (1988).
14. P.G. Estabrooks, *et al.*, Nucl. Phys. **B133**, 490 (1978).
15. A. Abele, *et al.*, Phys. Rev. **D57**, 3860 (1998).
16. V. Bernard, N. Kaiser, and U.-G. Meißner, Phys. Rev. **D43**, 2757 (1991).
17. S.N. Cherry and M.R. Pennington, Nucl. Phys. **A688**, 823 (2001).
18. J.M. Link, *et al.*, Phys. Lett. **B648**, 156 (2007).
19. M. Ablikim, *et al.*, Phys. Lett. **B633**, 681 (2006).
20. F.K. Guo, *et al.*, Nucl. Phys. **A773**, 78 (2006).
21. D. Epifanov, *et al.*, Phys. Lett. **B654**, 65 (2007).
22. C. Cawlfeld, *et al.*, Phys. Rev. **D74**, 031108R (2006).
23. A.V. Anisovich and A.V. Sarantsev, Phys. Lett. **B413**, 137 (1997).
24. R. Delbourgo, *et al.*, Int. J. Mod. Phys. **A13**, 657 (1998).
25. J.A. Oller, *et al.*, Phys. Rev. **D60**, 099906E (1999).
26. J.A. Oller and E. Oset, Phys. Rev. **D60**, 074023 (1999).
27. C.M. Shakin and H. Wang, Phys. Rev. **D63**, 014019 (2001).
28. M.D. Scadron, *et al.*, Nucl. Phys. **A724**, 391 (2003).
29. D.V. Bugg, Phys. Lett. **B572**, 1 (2003).
30. M. Ishida, Prog. Theor. Phys. Supp. **149**, 190 (2003).
31. H.Q. Zheng, *et al.*, Nucl. Phys. **A733**, 235 (2004).
32. Z.Y. Zhou and H.Q. Zheng, Nucl. Phys. **A775**, 212 (2006).
33. J.M. Link, *et al.*, Phys. Lett. **B653**, 1 (2007).
34. B. Aubert, *et al.*, Phys. Rev. **D76**, 011102R (2007).
35. S. Kopp, *et al.*, Phys. Rev. **D63**, 092001 (2001).
36. J.M. Link, *et al.*, Phys. Lett. **B535**, 43 (2002).
37. J.M. Link, *et al.*, Phys. Lett. **B621**, 72 (2005).
38. M. Jamin, *et al.*, Nucl. Phys. **B587**, 331 (2000).
39. D. Black, Phys. Rev. **D64**, 014031 (2001).
40. J.A. Oller, Nucl. Phys. **A727**, 353 (2003).
41. E. Van Beveren, *et al.*, Z. Phys. **C30**, 615 (1986).
42. J.R. Pelaez, Mod. Phys. Lett. **A19**, 2879 (2004).
43. S. Descotes-Genon and B. Moussallam, Eur. Phys. J. **C48**, 553 (2006).
44. J.J. Dudek *et al.*, Phys. Rev. Lett. **113**, 182001 (2014).
45. M. Bargiotti, *et al.*, Eur. Phys. J. **C26**, 371 (2003).
46. D. Barberis, *et al.*, Phys. Lett. **B440**, 225 (1998).
47. M.J. Corden, *et al.*, Nucl. Phys. **B144**, 253 (1978).
48. C. Defoix, *et al.*, Nucl. Phys. **B44**, 125 (1972).
49. Z. Bai, *et al.*, Phys. Rev. Lett. **65**, 2507 (1990).
50. T. Bolton, *et al.*, Phys. Rev. Lett. **69**, 1328 (1992).
51. C. Amsler, *et al.*, Phys. Lett. **B353**, 571 (1995).
52. S.M. Flatte, Phys. Lett. **63B**, 224 (1976).
53. C. Amsler, *et al.*, Phys. Lett. **B333**, 277 (1994).
54. G. Janssen, *et al.*, Phys. Rev. **D52**, 2690 (1995).
55. D.V. Bugg, Phys. Rev. **D78**, 074023 (2008).
56. P. Rubin, *et al.*, Phys. Rev. **D78**, 072003 (2008).
57. A.D. Martin and E.N. Ozmutlu, Nucl. Phys. **B158**, 520 (1979).
58. S.D. Protopopescu, *et al.*, Phys. Rev. **D7**, 1279 (1973).
59. G. Grayer, *et al.*, Nucl. Phys. **B75**, 189 (1974).
60. H. Becker, *et al.*, Nucl. Phys. **B151**, 46 (1979).
61. L. Rosselet, *et al.*, Phys. Rev. **D15**, 574 (1977).
62. S. Pislak, *et al.*, Phys. Rev. Lett. **87**, 221801 (2001).
63. J.R. Batley, *et al.*, Eur. Phys. J. **C70**, 635 (2010).
64. W. Wetzel, *et al.*, Nucl. Phys. **B115**, 208 (1976).
65. V.A. Polychronakos, *et al.*, Phys. Rev. **D19**, 1317 (1979).
66. D. Cohen, *et al.*, Phys. Rev. **D22**, 2595 (1980).
67. A. Etkin, *et al.*, Phys. Rev. **D25**, 1786 (1982).
68. C. Amsler, *et al.*, Phys. Lett. **B342**, 433 (1995).
69. C. Amsler, *et al.*, Phys. Lett. **B355**, 425 (1995).
70. A. Abele, *et al.*, Phys. Lett. **B380**, 453 (1996).
71. D.M. Alde, *et al.*, Phys. Lett. **B397**, 250 (1997).
72. R. Kaminski, L. Lesniak, and K. Rybicki, Z. Phys. **C74**, 79 (1997).
73. E.M. Aitala, *et al.*, Phys. Rev. Lett. **86**, 770 (2001).
74. J.M. Link, *et al.*, Phys. Lett. **B585**, 200 (2004).
75. G. Bonvicini, *et al.*, Phys. Rev. **D76**, 012001 (2007).
76. J.E. Augustin and G. Cosme, Nucl. Phys. **B320**, 1 (1989).
77. M. Ablikim, *et al.*, Phys. Lett. **B598**, 149 (2004).
78. A. Gallegos, *et al.*, Phys. Rev. **D69**, 074033 (2004).

79. M. Ablikim, *et al.*, Phys. Lett. **B645**, 19 (2007).
80. A. Dobado and J.R. Pelaez, Phys. Rev. **D56**, 3057 (1997).
81. I. Caprini, Phys. Rev. **D77**, 114019 (2008).
82. R. Garcia-Martin, J.R. Pelaez, and F.J. Yndurain, Phys. Rev. **D76**, 074034 (2007).
83. V.V. Anisovich, *et al.*, Sov. Phys. Usp. **41**, 419 (1998).
84. V.V. Anisovich, Int. J. Mod. Phys. **A21**, 3615 (2006).
85. B.S. Zou and D.V. Bugg, Phys. Rev. **D48**, R3948 (1993).
86. N.A. Tornqvist and M. Roos, Phys. Rev. Lett. **76**, 1575 (1996).
87. B.S. Zou and D.V. Bugg, Phys. Rev. **D50**, 3145 (1994).
88. N.N. Achasov and G.N. Shestakov, Phys. Rev. **D49**, 5779 (1994).
89. M.P. Locher, *et al.*, Eur. Phys. J. **C4**, 317 (1998).
90. J.A. Oller and E. Oset, Nucl. Phys. **A652**, 407 (1999).
91. T. Hannah, Phys. Rev. **D60**, 017502 (1999).
92. R. Kaminski, *et al.*, Phys. Rev. **D50**, 3145 (1994).
93. R. Kaminski, *et al.*, Phys. Lett. **B413**, 130 (1997).
94. R. Kaminski, *et al.*, Eur. Phys. J. **C9**, 141 (1999).
95. M. Ishida, *et al.*, Prog. Theor. Phys. **104**, 203 (2000).
96. Y.S. Surovtsev, *et al.*, Phys. Rev. **D61**, 054024 (2001).
97. M. Ishida, *et al.*, Phys. Lett. **B518**, 47 (2001).
98. M. Ablikim, *et al.*, Phys. Lett. **B598**, 149 (2004).
99. Z.Y. Zhou, *et al.*, JHEP **0502**, 043 (2005).
100. D.V. Bugg, *et al.*, J. Phys. **G34**, 151 (2007).
101. G. Bonvicini, *et al.*, Phys. Rev. **D76**, 012001 (2007).
102. D. Morgan and M.R. Pennington, Z. Phys. **C48**, 623 (1990).
103. M.R. Pennington, Phys. Rev. Lett. **97**, 011601 (2006).
104. M.R. Pennington, Mod. Phys. Lett. **A22**, 1439 (2007).
105. G. Mennessier, S. Narison, and W. Ochs, Phys. Lett. **B665**, 205 (2008).
106. R. Garcia-Martin and B. Moussallam, Eur. Phys. J. **C70**, 155 (2010).
107. M. Hoferichter, *et al.*, Eur. Phys. J. **C71**, 1743 (2011).
108. L.Y. Dai and M.R. Pennington, Phys. Rev. **D90**, 036004 (2014).
109. J.R. Pelaez and G. Rios, Phys. Rev. Lett. **97**, 242002 (2006).
110. H.-X. Chen, A. Hosaka, and S.-L. Zhu, Phys. Lett. **B650**, 369 (2007).
111. F. Giacosa, Phys. Rev. **D75**, 054007 (2007).
112. L. Maiani, *et al.*, Eur. Phys. J. **C50**, 609 (2007).
113. N.N. Achasov and G.N. Shestakov, Phys. Rev. **D58**, 054011 (1998).
114. N.N. Achasov, *et al.*, Phys. Lett. **B479**, 53 (2000).
115. N.N. Achasov, *et al.*, Phys. Lett. **B485**, 349 (2000).
116. R.R. Akhmetshin, *et al.*, Phys. Lett. **B462**, 371 (1999).
117. A. Aloisio, *et al.*, Phys. Lett. **B536**, 209 (2002).
118. F. Ambrosino, *et al.*, Eur. Phys. J. **C49**, 473 (2007).
119. M. Boglione and M.R. Pennington, Eur. Phys. J. **C9**, 11 (1999).
120. T. Mori, *et al.*, Phys. Rev. **D75**, 051101R (2007).
121. N.N. Achasov and G.N. Shestakov, Phys. Rev. **D77**, 074020 (2008).
122. M.R. Pennington, *et al.*, Eur. Phys. J. **C56**, 1 (2008).
123. R. Fleischer, *et al.*, Eur. Phys. J. **C71**, 1832 (2011).
124. S. Stone and L. Zhang, Phys. Rev. Lett. **111**, 062001 (2013).
125. R. Aaij, *et al.*, (LHCb Collaboration), Phys. Rev. **D90**, 012003 (2014).
126. R. Aaij, *et al.*, (LHCb Collaboration), Phys. Rev. **D89**, 092006 (2014).
127. J.T. Daub, C. Hanhart, and B. Kubis, JHEP **1602**, 009 (2016).
128. D. Barberis, *et al.*, Phys. Lett. **B462**, 462 (1999).
129. D. Barberis, *et al.*, Phys. Lett. **B479**, 59 (2000).
130. M. Gaspero, Nucl. Phys. **A562**, 407 (1993).
131. A. Adamo, *et al.*, Nucl. Phys. **A558**, 13C (1993).
132. C. Amsler, *et al.*, Phys. Lett. **B322**, 431 (1994).
133. A. Abele, *et al.*, Eur. Phys. J. **C19**, 667 (2001).
134. A. Abele, *et al.*, Eur. Phys. J. **C21**, 261 (2001).
135. D. Barberis, *et al.*, Phys. Lett. **B471**, 440 (2000).
136. R. Jaffe, Phys. Rev. **D15**, 267,281 (1977).
137. M. Alford and R.L. Jaffe, Nucl. Phys. **B578**, 367 (2000).
138. L. Maiani, *et al.*, Phys. Rev. Lett. **93**, 212002 (2004).
139. L. Maiani, A.D. Polosa, and V. Riquer, Phys. Lett. **B651**, 129 (2007).
140. G. 'tHooft, *et al.*, Phys. Lett. **B662**, 424 (2008).
141. J. Weinstein and N. Isgur, Phys. Rev. **D41**, 2236 (1990).
142. T. Barnes, Phys. Lett. **B165**, 434 (1985).
143. Z.P. Li, *et al.*, Phys. Rev. **D43**, 2161 (1991).
144. R. Delbourgo, D. Lui, and M. Scadron, Phys. Lett. **B446**, 332 (1999).
145. J.L. Lucio and M. Napsuciale, Phys. Lett. **B454**, 365 (1999).
146. C. Hanhart, *et al.*, Phys. Rev. **D75**, 074015 (2007).
147. R.H. Lemmer, Phys. Lett. **B650**, 152 (2007).
148. T. Branz, T. Gutsche, and V. Lyubovitskij, Eur. Phys. J. **A37**, 303 (2008).
149. A. Deandrea, *et al.*, Phys. Lett. **B502**, 79 (2001).
150. K.M. Ecklund, *et al.*, Phys. Rev. **D80**, 052009 (2010).
151. N.N. Achasov, V.N. Ivanchenko, Nucl. Phys. **B315**, 465 (1989).
152. J.A. Oller, *et al.*, Nucl. Phys. **A714**, 161 (2003).
153. Y.S. Kalashnikova, *et al.*, Eur. Phys. J. **A24**, 437 (2005).
154. Y.S. Kalashnikova, *et al.*, Phys. Rev. **D78**, 058501 (2008).
155. N.N. Achasov and A.V. Kiselev, Phys. Rev. **D78**, 058502 (2008).
156. M. Boglione and M.R. Pennington, Eur. Phys. J. **C30**, 503 (2003).
157. F. Giacosa and G. Pagliara, Phys. Rev. **C76**, 065204 (2007).
158. S. Narison, Nucl. Phys. **B96**, 244 (2001).
159. H.J. Lee, Eur. Phys. J. **A30**, 423 (2006).
160. H.X. Chen, A. Hosaka, and S.L. Zhu, Phys. Rev. **D76**, 094025 (2007).
161. J. Sugiyama, *et al.*, Phys. Rev. **D76**, 114010 (2007).
162. T. Kojo and D. Jido, Phys. Rev. **D78**, 114005 (2008).
163. D. Black, *et al.*, Phys. Rev. **D59**, 074026 (1999).
164. A.H. Fariborz *et al.*, Phys. Rev. **D90**, 033009 (2014).

## Meson Particle Listings

 $f_0(500)$ 

165. M. Scadron, *Eur. Phys. J.* **C6**, 141 (1999).  
 166. M. Ishida, *Prog. Theor. Phys.* **101**, 661 (1999).  
 167. N. Tornqvist, *Eur. Phys. J.* **C11**, 359 (1999).  
 168. M. Napsuciale and S. Rodriguez, *Phys. Lett.* **B603**, 195 (2004).  
 169. M. Napsuciale and S. Rodriguez, *Phys. Rev.* **D70**, 094043 (2004).  
 170. J.R. Pelaez, *Phys. Rev. Lett.* **92**, 102001 (2004).  
 171. J.T. Londergan et al., *Phys. Lett.* **B729**, 9 (2014).  
 172. S. Weinberg, *Phys. Rev.* **130**, 776 (1963).  
 173. D. Morgan and M.R. Pennington, *Phys. Lett. B* **258** (1991) 444 [*Phys. Lett. B* **269** (1991) 477].  
 174. D. Morgan, *Nucl. Phys.* **A543**, 632 (1992).  
 175. N. Tornqvist, *Phys. Rev.* **D51**, 5312 (1995).  
 176. V. Baru, *et al.*, *Phys. Lett.* **B586**, 53 (2004).  
 177. N.N. Achasov, *et al.*, *Phys. Lett.* **B88**, 367 (1979).  
 178. J.-J. Wu, *et al.*, *Phys. Rev.* **D75**, 114012 (2007).  
 179. C. Hanhart, *et al.*, *Phys. Rev.* **D76**, 074028 (2007).  
 180. M. Ablikim, *et al.*, *Phys. Rev.* **D83**, 032003 (2011).  
 181. L. Roca, *Phys. Rev.* **D88**, 014045 (2013).  
 182. N.A. Tornqvist, *Z. Phys.* **C68**, 647 (1995).  
 183. E. Van Beveren, *Eur. Phys. J.* **C22**, 493 (2001).  
 184. M. Boglione and M.R. Pennington, *Phys. Rev.* **D65**, 114010 (2002).  
 185. A. Abele, *et al.*, *Phys. Lett.* **B385**, 425 (1996).  
 186. D. Black, *et al.*, *Phys. Rev.* **D61**, 074001 (2000).  
 187. C. Amsler, *et al.*, *Phys. Lett.* **B639**, 165 (2006).  
 188. M. Acciarri, *et al.*, *Phys. Lett.* **B501**, 173 (2001).  
 189. R. Barate, *et al.*, *Phys. Lett.* **B472**, 189 (2000).  
 190. S. Uehara, *et al.*, *Phys. Rev.* **D78**, 052004 (2008).  
 191. C. Amsler, *Phys. Lett.* **B541**, 22 (2002).  
 192. S. Uehara, *et al.*, *Prog. Theor. Exp. Phys.* **2013**, 123C01 (2013).  
 193. F.E. Close and Q. Zhao, *Phys. Rev.* **D71**, 094022 (2005).  
 194. S. Uehara, *et al.*, *Phys. Rev.* **D80**, 032001 (2009).  
 195. F.E. Close, *et al.*, *Phys. Lett.* **B397**, 333 (1997).  
 196. F.E. Close, *Phys. Lett.* **B419**, 387 (1998).  
 197. A. Kirk, *Phys. Lett.* **B489**, 29 (2000).  
 198. M. Ablikim, *et al.*, *Phys. Lett.* **B603**, 138 (2004).  
 199. M. Ablikim, *et al.*, *Phys. Lett.* **B607**, 243 (2005).  
 200. C. Amsler and F.E. Close, *Phys. Rev.* **D53**, 295 (1996).  
 201. F.E. Close and A. Kirk, *Eur. Phys. J.* **C21**, 531 (2001).  
 202. S. Janowski, *Phys. Rev.* **D90**, 114005 (2014).  
 203. P. Minkowski and W. Ochs, *Eur. Phys. J.* **C9**, 283 (1999).  
 204. W. Lee and D. Weingarten, *Phys. Rev.* **D61**, 014015 (2000).  
 205. M. Chanowitz, *Phys. Rev. Lett.* **95**, 172001 (2005).  
 206. M. Albaladejo and J.A. Oller, *Phys. Rev. Lett.* **101**, 252002 (2008).  
 207. P. Minkowski, W. Ochs, *Eur. Phys. J.* **C39**, 71 (2005).  
 208. A. Garmash, *et al.*, *Phys. Rev.* **D71**, 092003 (2005).  
 209. B. Aubert, *et al.*, *Phys. Rev.* **D74**, 032003 (2006).  
 210. B. Aubert, *et al.*, *Phys. Rev. Lett.* **99**, 161802 (2007).  
 211. M. Ablikim, *et al.*, *Phys. Lett.* **B642**, 441 (2006).  
 212. R. Molina, *et al.*, *Phys. Rev.* **D78**, 114018 (2008).

213. C. Garcia-Recio, *et al.*, *Phys. Rev.* **D87**, 096006 (2013).  
 214. L.S. Geng, *et al.*, *Eur. Phys. J.* **A44**, 305 (2010).  
 215. T. Branz, *et al.*, *Phys. Rev.* **D81**, 054037 (2010).  
 216. A. Martinez Torres, *et al.*, *Phys. Lett.* **B719**, 388 (2013).  
 217. M. Ablikim, *et al.*, *Phys. Rev. Lett.* **96**, 162002 (2006).

 $f_0(500)$  T-MATRIX POLE  $\sqrt{s}$ Note that  $\Gamma \approx 2 \operatorname{Im}(\sqrt{s}_{\text{pole}})$ .

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>(400–550)–<math>i</math>(200–350) OUR ESTIMATE</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
(440 $\pm$ 10)– $i$ (238 $\pm$ 10)	<sup>1</sup> ALBALADEJO 12	RVUE	Compilation
(445 $\pm$ 25)– $i$ (278 $\pm$ <sup>22</sup> <sub>18</sub> )	<sup>2,3</sup> GARCIA-MAR..11	RVUE	Compilation
(457 $\pm$ <sup>14</sup> <sub>13</sub> )– $i$ (279 $\pm$ <sup>11</sup> <sub>7</sub> )	<sup>2,4</sup> GARCIA-MAR..11	RVUE	Compilation
(442 $\pm$ <sup>5</sup> <sub>8</sub> )– $i$ (274 $\pm$ <sup>6</sup> <sub>5</sub> )	<sup>5</sup> MOUSSALLAM11	RVUE	Compilation
(452 $\pm$ 13)– $i$ (259 $\pm$ 16)	<sup>6</sup> MENNESSIER 10	RVUE	Compilation
(448 $\pm$ 43)– $i$ (266 $\pm$ 43)	<sup>7</sup> MENNESSIER 10	RVUE	Compilation
(455 $\pm$ <sup>6</sup> <sub>13</sub> )– $i$ (278 $\pm$ <sup>6</sup> <sub>43</sub> )	<sup>8</sup> CAPRINI 08	RVUE	Compilation
(463 $\pm$ <sup>6</sup> <sub>17</sub> )– $i$ (259 $\pm$ <sup>6</sup> <sub>34</sub> )	<sup>9</sup> CAPRINI 08	RVUE	Compilation
(552 $\pm$ <sup>84</sup> <sub>106</sub> )– $i$ (232 $\pm$ <sup>81</sup> <sub>72</sub> )	<sup>10</sup> ABLIKIM 07A	BES2	$\psi(2S) \rightarrow \pi^+ \pi^- J/\psi$
(466 $\pm$ 18)– $i$ (223 $\pm$ 28)	<sup>11</sup> BONVICINI 07	CLEO	$D^+ \rightarrow \pi^- \pi^+ \pi^+$
(472 $\pm$ 30)– $i$ (271 $\pm$ 30)	<sup>12</sup> BUGG 07A	RVUE	Compilation
(484 $\pm$ 17)– $i$ (255 $\pm$ 10)	GARCIA-MAR..07	RVUE	Compilation
(430)– $i$ (325)	<sup>13</sup> ANISOVICH 06	RVUE	Compilation
(441 $\pm$ <sup>16</sup> <sub>8</sub> )– $i$ (272 $\pm$ <sup>9</sup> <sub>12.5</sub> )	<sup>14</sup> CAPRINI 06	RVUE	$\pi\pi \rightarrow \pi\pi$
(470 $\pm$ 50)– $i$ (285 $\pm$ 25)	<sup>15</sup> ZHOU 05	RVUE	
(541 $\pm$ 39)– $i$ (252 $\pm$ 42)	<sup>16</sup> ABLIKIM 04A	BES2	$J/\psi \rightarrow \omega \pi^+ \pi^-$
(528 $\pm$ 32)– $i$ (207 $\pm$ 23)	<sup>17</sup> GALLEGOS 04	RVUE	Compilation
(440 $\pm$ 8)– $i$ (212 $\pm$ 15)	<sup>18</sup> PELAEZ 04A	RVUE	$\pi\pi \rightarrow \pi\pi$
(533 $\pm$ 25)– $i$ (249 $\pm$ 25)	<sup>19</sup> BUGG 03	RVUE	
517– $i$ 240	BLACK 01	RVUE	$\pi^0 \pi^0 \rightarrow \pi^0 \pi^0$
(470 $\pm$ 30)– $i$ (295 $\pm$ 20)	<sup>14</sup> COLANGELO 01	RVUE	$\pi\pi \rightarrow \pi\pi$
(535 $\pm$ <sup>48</sup> <sub>36</sub> )– $i$ (155 $\pm$ <sup>76</sup> <sub>53</sub> )	<sup>20</sup> SHIDA 01		$\Upsilon(3S) \rightarrow \Upsilon \pi\pi$
610 $\pm$ 14– $i$ 2620 $\pm$ 26	<sup>21</sup> SUROVTSEV 01	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
(540 $\pm$ <sup>36</sup> <sub>29</sub> )– $i$ (193 $\pm$ <sup>32</sup> <sub>40</sub> )	ISHIDA 00B		$\rho\bar{\rho} \rightarrow \pi^0 \pi^0 \pi^0$
445– $i$ 235	HANNAH 99	RVUE	$\pi$ scalar form factor
(523 $\pm$ 12)– $i$ (259 $\pm$ 7)	KAMINSKI 99	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, \sigma\sigma$
442– $i$ 227	OLLER 99	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
469– $i$ 203	OLLER 99B	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
445– $i$ 221	OLLER 99C	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, \eta\eta$
(1530 $\pm$ <sup>90</sup> <sub>250</sub> )– $i$ (560 $\pm$ 40)	ANISOVICH 98B	RVUE	Compilation
420– $i$ 212	LOCHER 98	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
440– $i$ 245	<sup>22</sup> DOBADO 97	RVUE	Compilation
(602 $\pm$ 26)– $i$ (196 $\pm$ 27)	<sup>23</sup> SHIDA 97		$\pi\pi \rightarrow \pi\pi$
(537 $\pm$ 20)– $i$ (250 $\pm$ 17)	<sup>24</sup> KAMINSKI 97B	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, 4\pi$
470– $i$ 250	<sup>25,26</sup> TORNQVIST 96	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, K\pi, \eta\pi$
387– $i$ 305	<sup>26,27</sup> JANSSEN 95	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
420– $i$ 370	<sup>28</sup> ACHASOV 94	RVUE	$\pi\pi \rightarrow \pi\pi$
(506 $\pm$ 10)– $i$ (247 $\pm$ 3)	KAMINSKI 94	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
370– $i$ 356	<sup>29</sup> ZOU 94B	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
408– $i$ 342	<sup>26,29</sup> ZOU 93	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
470– $i$ 208	<sup>30</sup> VANBEVEREN 86	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, \eta\eta,$
(750 $\pm$ 50)– $i$ (450 $\pm$ 50)	<sup>31</sup> ESTABROOKS 79	RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
(660 $\pm$ 100)– $i$ (320 $\pm$ 70)	PROTOPOP... 73	HBC	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
650– $i$ 370	<sup>32</sup> BASDEVANT 72	RVUE	$\pi\pi \rightarrow \pi\pi$

<sup>1</sup> Applying the chiral unitary approach at NLO to the  $K_{e4}$  data of BATLEY 10 and  $\pi N \rightarrow \pi\pi N$  data of HYAMS 73, GRAYER 74, and PROTOPODESCU 73.

<sup>2</sup> Uses the  $K_{e4}$  data of BATLEY 10c and the  $\pi N \rightarrow \pi\pi N$  data of HYAMS 73, GRAYER 74, and PROTOPODESCU 73.

<sup>3</sup> Analytic continuation using Roy equations.

<sup>4</sup> Analytic continuation using GKPY equations.

<sup>5</sup> Using Roy equations.

<sup>6</sup> Average of three variants of the analytic K-matrix model. Uses the  $K_{e4}$  data of BATLEY 08A and the  $\pi N \rightarrow \pi\pi N$  data of HYAMS 73 and GRAYER 74.

<sup>7</sup> Average of the analyses of three data sets in the K-matrix model. Uses the data of BATLEY 08A, HYAMS 73, and GRAYER 74, partially of COHEN 80 or ETKIN 82B.

<sup>8</sup> From the  $K_{e4}$  data of BATLEY 08A and  $\pi N \rightarrow \pi\pi N$  data of HYAMS 73.

<sup>9</sup> From the  $K_{e4}$  data of BATLEY 08A and  $\pi N \rightarrow \pi\pi N$  data of PROTOPODESCU 73, GRAYER 74, and ESTABROOKS 74.

<sup>10</sup> From a mean of three different  $f_0(500)$  parametrizations. Uses 40k events.

<sup>11</sup> From an isobar model using 2.6k events.

<sup>12</sup> Reanalysis of ABLIKIM 04A, PISLAK 01, and HYAMS 73 data.

<sup>13</sup> Using the N/D method.

<sup>14</sup> From the solution of the Roy equation (ROY 71) for the isoscalar S-wave and using a phase-shift analysis of HYAMS 73 and PROTOPODESCU 73 data.

See key on page 601

## Meson Particle Listings

 $f_0(500)$ 

- 15 Reanalysis of the data from PROTOPODESCU 73, ESTABROOKS 74, GRAYER 74, ROSSELET 77, PISLAK 03, and AKHMETSHIN 04.  
 16 From a mean of six different analyses and  $f_0(500)$  parameterizations.  
 17 Using data on  $\psi(2S) \rightarrow J/\psi \pi \pi$  from BAI 00E and on  $\Upsilon(nS) \rightarrow \Upsilon(mS) \pi \pi$  from BUTLER 94B and ALEXANDER 98.  
 18 Reanalysis of data from PROTOPODESCU 73, ESTABROOKS 74, GRAYER 74, and COHEN 80 in the unitarized ChPT model.  
 19 From a combined analysis of HYAMS 73, AUGUSTIN 89, AITALA 01B, and PISLAK 01.  
 20 A similar analysis (KOMADA 01) finds  $(580^{+79}_{-30}) - i(190^{+107}_{-49})$  MeV.  
 21 Coupled channel reanalysis of BATON 70, BENSINGER 71, BAILLON 72, HYAMS 73, HYAMS 75, ROSSELET 77, COHEN 80, and ETKIN 82B using the uniformizing variable.  
 22 Using the inverse amplitude method and data of ESTABROOKS 73, GRAYER 74, and PROTOPODESCU 73.  
 23 Reanalysis of data from HYAMS 73, GRAYER 74, SRINIVASAN 75, and ROSSELET 77 using the interfering amplitude method.  
 24 Average and spread of 4 variants ("up" and "down") of KAMINSKI 97B 3-channel model.  
 25 Uses data from BEIER 72B, OCHS 73, HYAMS 73, GRAYER 74, ROSSELET 77, CASON 83, ASTON 88, and ARMSTRONG 91B. Coupled channel analysis with flavor symmetry and all light two-pseudoscalars systems.  
 26 Demonstrates explicitly that  $f_0(500)$  and  $f_0(1370)$  are two different poles.  
 27 Analysis of data from FALVARD 88.  
 28 Analysis of data from OCHS 73, ESTABROOKS 75, ROSSELET 77, and MUKHIN 80.  
 29 Analysis of data from OCHS 73, GRAYER 74, and ROSSELET 77.  
 30 Coupled-channel analysis using data from PROTOPODESCU 73, HYAMS 73, HYAMS 75, GRAYER 74, ESTABROOKS 74, ESTABROOKS 75, FROGGATT 77, CORDEN 79, BISWAS 81.  
 31 Analysis of data from APEL 73, GRAYER 74, CASON 76, PAWLICKI 77. Includes spread and errors of 4 solutions.  
 32 Analysis of data from BATON 70, BENSINGER 71, COLTON 71, BAILLON 72, PROTOPODESCU 73, and WALKER 67.

 $f_0(500)$  BREIT-WIGNER MASS OR K-MATRIX POLE PARAMETERS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>(400-550) OUR ESTIMATE</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$513 \pm 32$	33 MURAMATSU 02	CLEO	$e^+ e^- \approx 10$ GeV
$478^{+24}_{-23} \pm 17$	AITALA 01B	E791	$D^+ \rightarrow \pi^- \pi^+ \pi^+$
$563^{+58}_{-29}$	34 ISHIDA 01		$\Upsilon(3S) \rightarrow \Upsilon \pi \pi$
555	35 ASNER 00	CLE2	$\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
$540 \pm 36$	ISHIDA 00B		$p\bar{p} \rightarrow \pi^0 \pi^0 \pi^0$
$750 \pm 4$	ALEKSEEV 99	SPEC	$1.78 \pi^- p_{\text{polar}} \rightarrow \pi^- \pi^+ n$
$744 \pm 5$	ALEKSEEV 98	SPEC	$1.78 \pi^- p_{\text{polar}} \rightarrow \pi^- \pi^+ n$
$759 \pm 5$	36 TROYAN 98		$5.2 n p \rightarrow n p \pi^+ \pi^-$
$780 \pm 30$	ALDE 97	GAM2	$450 p p \rightarrow p p \pi^0 \pi^0$
$585 \pm 20$	37 ISHIDA 97		$\pi \pi \rightarrow \pi \pi$
$761 \pm 12$	38 SVEC 96	RVUE	$6-17 \pi N_{\text{polar}} \rightarrow \pi^+ \pi^- N$
$\sim 860$	39,40 TORNQVIST 96	RVUE	$\pi \pi \rightarrow \pi \pi, K \bar{K}, K \pi, \eta \pi$
$1165 \pm 50$	41,42 ANISOVICH 95	RVUE	$\pi^- p \rightarrow \pi^0 \pi^0 n, \bar{p} p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \pi^0 \eta, \pi^0 \eta \eta$
$\sim 1000$	43 ACHASOV 94	RVUE	$\pi \pi \rightarrow \pi \pi$
$414 \pm 20$	38 AUGUSTIN 89	DM2	

- 33 Statistical uncertainty only.  
 34 A similar analysis (KOMADA 01) finds  $526^{+48}_{-37}$  MeV.  
 35 From the best fit of the Dalitz plot.  
 36  $6\sigma$  effect, no PWA.  
 37 Reanalysis of data from HYAMS 73, GRAYER 74, SRINIVASAN 75, and ROSSELET 77 using the interfering amplitude method.  
 38 Breit-Wigner fit to S-wave intensity measured in  $\pi N \rightarrow \pi^- \pi^+ N$  on polarized targets. The fit does not include  $f_0(980)$ .  
 39 Uses data from ASTON 88, OCHS 73, HYAMS 73, ARMSTRONG 91B, GRAYER 74, CASON 83, ROSSELET 77, and BEIER 72B. Coupled channel analysis with flavor symmetry and all light two-pseudoscalars systems.  
 40 Also observed by ASNER 00 in  $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$  decays.  
 41 Uses  $\pi^0 \pi^0$  data from ANISOVICH 94, AMSLER 94D, and ALDE 95B,  $\pi^+ \pi^-$  data from OCHS 73, GRAYER 74 and ROSSELET 77, and  $\eta \eta$  data from ANISOVICH 94.  
 42 The pole is on Sheet III. Demonstrates explicitly that  $f_0(500)$  and  $f_0(1370)$  are two different poles.  
 43 Analysis of data from OCHS 73, ESTABROOKS 75, ROSSELET 77, and MUKHIN 80.

 $f_0(500)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>(400-700) OUR ESTIMATE</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$335 \pm 67$	44 MURAMATSU 02	CLEO	$e^+ e^- \approx 10$ GeV
$324^{+42}_{-40} \pm 21$	AITALA 01B	E791	$D^+ \rightarrow \pi^- \pi^+ \pi^+$
$372^{+229}_{-95}$	45 ISHIDA 01		$\Upsilon(3S) \rightarrow \Upsilon \pi \pi$
540	46 ASNER 00	CLE2	$\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
$372 \pm 80$	ISHIDA 00B		$p\bar{p} \rightarrow \pi^0 \pi^0 \pi^0$
$119 \pm 13$	ALEKSEEV 99	SPEC	$1.78 \pi^- p_{\text{polar}} \rightarrow \pi^- \pi^+ n$
$77 \pm 22$	ALEKSEEV 98	SPEC	$1.78 \pi^- p_{\text{polar}} \rightarrow \pi^- \pi^+ n$

$35 \pm 12$	47 TROYAN 98		$5.2 n p \rightarrow n p \pi^+ \pi^-$
$780 \pm 60$	ALDE 97	GAM2	$450 p p \rightarrow p p \pi^0 \pi^0$
$385 \pm 70$	48 ISHIDA 97		$\pi \pi \rightarrow \pi \pi$
$290 \pm 54$	49 SVEC 96	RVUE	$6-17 \pi N_{\text{polar}} \rightarrow \pi^+ \pi^- N$
$\sim 880$	50,51 TORNQVIST 96	RVUE	$\pi \pi \rightarrow \pi \pi, K \bar{K}, K \pi, \eta \pi$
$460 \pm 40$	52,53 ANISOVICH 95	RVUE	$\pi^- p \rightarrow \pi^0 \pi^0 n, \bar{p} p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \pi^0 \eta, \pi^0 \eta \eta$
$\sim 3200$	54 ACHASOV 94	RVUE	$\pi \pi \rightarrow \pi \pi$
$494 \pm 58$	49 AUGUSTIN 89	DM2	

- 44 Statistical uncertainty only.  
 45 A similar analysis (KOMADA 01) finds  $301^{+145}_{-100}$  MeV.  
 46 From the best fit of the Dalitz plot.  
 47  $6\sigma$  effect, no PWA.  
 48 Reanalysis of data from HYAMS 73, GRAYER 74, SRINIVASAN 75, and ROSSELET 77 using the interfering amplitude method.  
 49 Breit-Wigner fit to S-wave intensity measured in  $\pi N \rightarrow \pi^- \pi^+ N$  on polarized targets. The fit does not include  $f_0(980)$ .  
 50 Uses data from ASTON 88, OCHS 73, HYAMS 73, ARMSTRONG 91B, GRAYER 74, CASON 83, ROSSELET 77, and BEIER 72B. Coupled channel analysis with flavor symmetry and all light two-pseudoscalars systems.  
 51 Also observed by ASNER 00 in  $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$  decays.  
 52 Uses  $\pi^0 \pi^0$  data from ANISOVICH 94, AMSLER 94D, and ALDE 95B,  $\pi^+ \pi^-$  data from OCHS 73, GRAYER 74 and ROSSELET 77, and  $\eta \eta$  data from ANISOVICH 94.  
 53 The pole is on Sheet III. Demonstrates explicitly that  $f_0(500)$  and  $f_0(1370)$  are two different poles.  
 54 Analysis of data from OCHS 73, ESTABROOKS 75, ROSSELET 77, and MUKHIN 80.

 $f_0(500)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \pi \pi$	dominant
$\Gamma_2 \quad \gamma \gamma$	seen

 $f_0(500)$  PARTIAL WIDTHS

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.05 \pm 0.21$	55 DAI 14A	RVUE	Compilation
$1.7 \pm 0.4$	56 HOFERICHTER11	RVUE	Compilation
$3.08 \pm 0.82$	57 MENNESSIER 11	RVUE	Compilation
$2.08 \pm 0.2 \pm 0.07_{-0.04}$	58 MOUSSALLAM11	RVUE	Compilation
2.08	59 MAO 09	RVUE	Compilation
$1.2 \pm 0.4$	60 BERNABEU 08	RVUE	
$3.9 \pm 0.6$	57 MENNESSIER 08	RVUE	$\gamma \gamma \rightarrow \pi^+ \pi^-, \pi^0 \pi^0$
$1.8 \pm 0.4$	61 OLLER 08	RVUE	Compilation
$1.68 \pm 0.15$	61,62 OLLER 08A	RVUE	Compilation
$3.1 \pm 0.5$	63,64 PENNINGTON 08	RVUE	Compilation
$2.4 \pm 0.4$	64,65 PENNINGTON 08	RVUE	Compilation
$4.1 \pm 0.3$	66 PENNINGTON 06	RVUE	$\gamma \gamma \rightarrow \pi^0 \pi^0$
$3.8 \pm 1.5$	67,68 BOGLIONE 99	RVUE	$\gamma \gamma \rightarrow \pi^+ \pi^-, \pi^0 \pi^0$
$5.4 \pm 2.3$	67 MORGAN 90	RVUE	$\gamma \gamma \rightarrow \pi^+ \pi^-, \pi^0 \pi^0$
$10 \pm 6$	COURAU 86	DM1	$e^+ e^- \rightarrow \pi^+ \pi^- e^+ e^-$

- 55 Using dispersive analysis with phases from GARCIA-MARTIN 11A and BUETTAKER 04 as input.  
 56 Using Roy-Steiner equations with  $\pi \pi$  phase shifts from an update of COLANGELO 01 and from GARCIA-MARTIN 11A.  
 57 Using an analytic K-matrix model.  
 58 Using dispersion integral with phase input from Roy equations and data from MARSISKE 90, BOYER 90, BEHREND 92, UEHARA 08A, and MORI 07.  
 59 Used dispersion theory. The value quoted used the  $f_0(500)$  pole position of  $457 - i276$  MeV.  
 60 Using  $p, n$  polarizabilities from PDG 06 and fitting to  $\pi \pi$  phase motion from GARCIA-MARTIN 07 and  $\sigma$ -poles from GARCIA-MARTIN 07 and CAPRINI 06.  
 61 Using twice-subtracted dispersion integrals.  
 62 Supersedes OLLER 08.  
 63 Solution A (preferred solution based on  $\chi^2$ -analysis).  
 64 Dispersion theory based amplitude analysis of BOYER 90, MARSISKE 90, BEHREND 92, and MORI 07.  
 65 Solution B (worse than solution A; still acceptable when systematic uncertainties are included).  
 66 Using unitarity and the  $\sigma$  pole position from CAPRINI 06.  
 67 This width could equally well be assigned to the  $f_0(1370)$ . The authors analyse data from BOYER 90 and MARSISKE 90 and report strong correlation with  $\gamma \gamma$  width of  $f_2(1270)$ .  
 68 Supersedes MORGAN 90.

# Meson Particle Listings

$f_0(500)$ ,  $\rho(770)$

## $f_0(500)$ REFERENCES

DAI 14A PR D90 036004 L.-Y. Dai, M.R. Pennington (CEBAF)  
ALBALADEJO 12 PR D86 034003 M. Albaladejo, J.A. Oller (MURC)  
GARCIA-MAR. 11 PRL 107 072001 R. Garcia-Martin *et al.* (MADR, CRAC)  
GARCIA-MAR. 11A PR D83 074004 R. Garcia-Martin *et al.* (MADR, CRAC)  
HOERICHTER 11 EPJ C71 1743 M. Hoferichter, D.R. Phillips, C. Schat (BONN+)  
MENNESSIER 11 PL B696 40 G. Mennessier, S. Narison, X.-G. Wang  
MOUSSALLAM 11 EPJ C71 1814 B. Moussallam  
BATLEY 10 PL B686 101 J.R. Batley *et al.* (CERN NA48/2 Collab.)  
BATLEY 10C EPJ C70 635 J.R. Batley *et al.* (CERN NA48/2 Collab.)  
MENNESSIER 10 PL B688 59 G. Mennessier, S. Narison, X.-G. Wang  
MAO 09 PR D79 116008 Y. Mao *et al.*  
BATLEY 08A EPJ C54 411 J.R. Batley *et al.* (CERN NA48/2 Collab.)  
BERNABEU 08 PRL 100 241804 J. Bernabeu, J. Prades (IFIC, GRAN)  
CAPRINI 08 PR D77 114019 I. Caprini  
MENNESSIER 08 PL B665 205 G. Mennessier, S. Narison, W. Ochs  
OLLER 08 PL B659 201 J.A. Oller, L. Roca, C. Schat (MURC, UBA)  
OLLER 08A EPJ A37 15 J.A. Oller, L. Roca (MURC)  
PENNINGTON 08 EPJ C56 1 M.R. Pennington *et al.*  
UEHARA 08A PR D78 052004 S. Uehara *et al.* (BELLE Collab.)  
ABLIKIM 07A PL B645 19 M. Ablikim *et al.* (BES Collab.)  
BONVICINI 07 PR D76 012001 G. Bonvicini *et al.* (CLEO Collab.)  
BUGG 07A JP G34 151 D.V. Bugg *et al.*  
GARCIA-MAR... 07 PR D76 074034 R. Garcia-Martin, J.R. Pelaez, F.J. Yndurain (BELLE Collab.)  
MORI 07 PR D75 051101 T. Mori *et al.*  
ANISOVICH 06 UMP A21 3615 V.V. Anisovich  
CAPRINI 06 PRL 96 132001 I. Caprini, G. Colangelo, H. Leutwyler (BCIP+)  
PDG 06 JP G33 1 W.-M. Yao *et al.* (PDG Collab.)  
PENNINGTON 06 PRL 97 011601 M.R. Pennington  
ZHOU 05 JHEP 0502 043 Z.Y. Zhou *et al.*  
ABLIKIM 04A PL B598 149 M. Ablikim *et al.* (BES Collab.)  
AKHMETSHIN 04 PL B578 285 R.R. Akhmetshin *et al.* (Novosibirsk CMD-2 Collab.)  
BUETTIGER 04 EPJ C33 409 P. Buettiker, S. Descotes-Genon, B. Moussallam  
GALLEGOS 04 PR D69 074033 A. Gallegos *et al.*  
PELAEZ 04A MPL A19 2879 J.R. Pelaez  
BUGG 03 PL B572 1 D.V. Bugg  
PISLAK 03 PR D67 072004 S. Pislak *et al.* (BNL E865 Collab.)  
Also S. Pislak *et al.* (BNL E865 Collab.)  
MURAMATSU 02 PRL 89 251802 H. Muramatsu *et al.* (CLEO Collab.)  
Also PRL 90 059901 (errata) E. Muramatsu *et al.* (CLEO Collab.)  
AITALA 01B PRL 86 770 M. Aitala *et al.* (FNAL E791 Collab.)  
BLACK 01 PR D64 014031 D. Black *et al.*  
COLANGELO 01 NP B603 125 G. Colangelo, J. Gasser, H. Leutwyler  
ISHIDA 01 PL B518 47 M. Ishida *et al.*  
KOMADA 01 PL B508 31 T. Komada *et al.*  
PISLAK 01 PRL 87 221801 S. Pislak *et al.* (BNL E865 Collab.)  
Also PR D67 072000 S. Pislak *et al.* (BNL E865 Collab.)  
Also PRL 105 019901E S. Pislak *et al.* (BNL E865 Collab.)  
SUROVTSSEV 01 PR D63 054024 Y.S. Surovtsev, D. Krupa, M. Nagy  
ASNER 00 PR D61 012002 D.M. Asner *et al.* (CLEO Collab.)  
BAI 00E PR D62 032002 J. Bai *et al.* (BES Collab.)  
ISHIDA 00B PTP 104 203 M. Ishida *et al.*  
ALEKSEEV 99 NP B541 3 I.G. Alekseev *et al.*  
BOGLIONE 99 EPJ C9 11 M. Boggione, M.R. Pennington  
HANNAH 99 PR D60 017502 T. Hannah  
KAMINSKI 99 EPJ C9 141 R. Kaminski, L. Lesniak, B. Loiseau (CRAC, PARIN)  
OLLER 99 PR D60 099906 (erratum) J.A. Oller *et al.*  
OLLER 99B NP A652 407 (erratum) J.A. Oller, E. Oset  
OLLER 99 PR D60 074023 J.A. Oller, E. Oset  
ALEKSEEV 98 PAN 61 174 I.G. Alekseev *et al.*  
ALEXANDER 98 PR D58 052004 J.P. Alexander *et al.* (CLEO Collab.)  
ANISOVICH 98B SPU 41 419 V.V. Anisovich *et al.*  
Translated from UFN 168 481.  
LOCHER 98 EPJ C4 317 M.P. Locher *et al.* (PSI)  
TROYAN 98 JINRRC 5-91 33 Y. Troyan *et al.*  
ALDE 97 PL B397 350 D.M. Alde *et al.* (GAMS Collab.)  
DOBADO 97 PR D56 3057 S. Dobado, J.R. Pelaez  
ISHIDA 97 PTP 98 1005 S. Ishida *et al.* (TOKY, MIYA, KEK)  
KAMINSKI 97B PL B413 130 R. Kaminski, L. Lesniak, B. Loiseau (CRAC, IPN)  
Also S. Ishida *et al.* (TOKY, MIYA, KEK)  
SVEC 96 PR D53 2343 M. Svec (MCGI)  
TORNVIST 96 PRL 76 1575 N.A. Tornqvist, M. Roos (HELSE)  
ALDE 95B ZPHY C66 375 D.M. Alde *et al.* (GAMS Collab.)  
ANISOVICH 95 PL B355 363 V.V. Anisovich *et al.* (PNPI, SERP)  
JANSSEN 95 PR D52 2690 G. Janssen *et al.* (STON, ADLD, JULI)  
ACHASOV 94 PR D49 5779 N.N. Achasov, G.N. Shestakov (NOVM)  
AMSLER 94D PL B333 277 C. Amisler *et al.* (Crystal Barrel Collab.)  
ANISOVICH 94 PL B323 233 V.V. Anisovich *et al.* (Crystal Barrel Collab.)  
BUTLER 94B PR D49 440 F. Butler *et al.* (CLEO Collab.)  
KAMINSKI 94 PR D50 3145 R. Kaminski, L. Lesniak, J.P. Maillet (CRAC+)  
ZOU 94B PR D50 591 B.S. Zou, D.V. Bugg (LOQM)  
ZOU 93 PR D48 R3948 B.S. Zou, D.V. Bugg (LOQM)  
BEHREND 92 ZPHY C56 381 H.J. Behrend (CELLO Collab.)  
ARMSTRONG 91B ZPHY C52 389 T.A. Armstrong *et al.* (ATHU, BARI, BIRM+)  
BOYER 90 PR D42 1350 J. Boyer *et al.* (Mark II Collab.)  
MARISKE 90 PR D41 3324 H. Mariske *et al.* (Crystal Ball Collab.)  
MORGAN 90 ZPHY C48 623 D. Morgan, M.R. Pennington (RAL, DURH)  
AUGUSTIN 89 NP B320 1 J.E. Augustin, G. Cosme (DM2 Collab.)  
ASTON 88 NP B296 493 D. Aston *et al.* (SLAC, NAGO, CINC, INUS)  
FALVARD 88 PR D38 2706 A. Falvard *et al.* (CLER, FRAS, LALO+)  
COURAU 86 NP B271 1 A. Courau *et al.* (CLER, LALO)  
VANBEVEREN 86 ZPHY C30 615 E. van Beveren *et al.* (NUM, BIEL)  
CASON 83 PR D28 1586 N.M. Cason *et al.* (NDAM, ANL)  
ETKIN 82B PR D25 1786 A. Etkin *et al.* (BNL, CUNY, TUFTS, VAND)  
BISWAS 81 PRL 47 1378 N.N. Biswas *et al.* (NDAM, ANL)  
COHEN 80 PR D22 2595 D. Cohen *et al.* (ANL) IUP  
MUKHIN 80 JETPL 32 601 K.N. Mukhin *et al.* (KIAE)  
Translated from ZETFP 32 616.  
CORDEN 79 NP B157 250 M.J. Corden *et al.* (BIRM, RHEL, TELA+) JP  
ESTABROOKS 79 PR D19 2678 P. Estabrooks (CARL)  
FROGGATT 77 NP B129 89 C.D. Froggatt, J.L. Petersen (GLAS, NORD)  
PAWLICKI 77 PR D15 3196 A.J. Pawlicki *et al.* (ANL) IJ  
ROSSELET 77 PR D15 574 L. Rosselet *et al.* (GEVA, SAEL)  
CASON 76 PRL 36 1485 N.M. Cason *et al.* (NDAM, ANL) IJ  
ESTABROOKS 75 NP B95 322 P.G. Estabrooks, A.D. Martin (DURH)  
HYAMS 75 NP B100 205 B.D. Hyams *et al.* (CERN, MPIM)  
SRINIVASAN 75 PR D12 681 V. Srinivasan *et al.* (NDAM, ANL)  
ESTABROOKS 74 NP B79 301 P.G. Estabrooks, A.D. Martin (DURH)  
GRAYEY 74 PR B75 189 G. Grayer *et al.* (CERN, MPIM)  
APEL 73 PL 41B 642 W.A. Apel *et al.* (KARL, PISA)  
ESTABROOKS 73 Tallahassee P.G. Estabrooks *et al.* (CERN, MPIM)  
HYAMS 73 NP B64 134 B.D. Hyams *et al.* (CERN, MPIM)  
OCHS 73 Thesis W. Ochs (MPIM, MUNI)  
PROTOPOPO... 73 PR D7 1279 S.D. Protopopescu *et al.* (LBL)  
BAILLON 72 PL 38B 555 P.H. Baillon *et al.* (SLAC)  
BASDEVANT 72 PL 41B 178 J.L. Basdevant, C.D. Froggatt, J.L. Petersen (CERN)  
BEIER 72B PRL 29 511 E.W. Beier *et al.* (PENN)  
BENSINGER 71 PL 36B 134 J.R. Bensinger *et al.* (WISC)  
COLTON 71 PR D3 2028 E.P. Colton *et al.* (LBL, FNAL, UCLA+)  
ROY 71 PL 36B 353 S.M. Roy

BATON 70 PL 33B 528  
WALKER 67 RMP 39 695

J.P. Baton, G. Laurens, J. Reigner  
W.D. Walker

(SACL)  
(WISC)

$\rho(770)$

$$I^G(J^{PC}) = 1^+(1^{--})$$

## THE $\rho(770)$

Updated May 2012 by S. Eidelman (Novosibirsk) and G. Venanzoni (Frascati).

The determination of the parameters of the  $\rho(770)$  is beset with many difficulties because of its large width. In physical region fits, the line shape does not correspond to a relativistic Breit-Wigner function with a  $P$ -wave width, but requires some additional shape parameter. This dependence on parameterization was demonstrated long ago [1]. Bose-Einstein correlations are another source of shifts in the  $\rho(770)$  line shape, particularly in multiparticle final state systems [2].

The same model-dependence afflicts any other source of resonance parameters, such as the energy-dependence of the phase shift  $\delta_1^+$ , or the pole position. It is, therefore, not surprising that a study of  $\rho(770)$  dominance in the decays of the  $\eta$  and  $\eta'$  reveals the need for specific dynamical effects, in addition to the  $\rho(770)$  pole [3,4].

The cleanest determination of the  $\rho(770)$  mass and width comes from  $e^+e^-$  annihilation and  $\tau$ -lepton decays. Analysis of ALEPH [5] showed that the charged  $\rho(770)$  parameters measured from  $\tau$ -lepton decays are consistent with those of the neutral one determined from  $e^+e^-$  data [6]. This conclusion is qualitatively supported by the later studies of CLEO [7] and Belle [8]. However, model-independent comparison of the two-pion mass spectrum in  $\tau$  decays, and the  $e^+e^- \rightarrow \pi^+\pi^-$  cross section, gave indications of discrepancies between the overall normalization:  $\tau$  data are about 3% higher than  $e^+e^-$  data [7,9]. A detailed analysis using such two-pion mass spectra from  $\tau$  decays measured by OPAL [10], CLEO [7], and ALEPH [11,12], as well as recent pion form factor measurements in  $e^+e^-$  annihilation by CMD-2 [13,14], showed that the discrepancy can be as high as 10% above the  $\rho$  meson [15,16]. This discrepancy remains after recent measurements of the two-pion cross section in  $e^+e^-$  annihilation at KLOE [17,18] and SND [19,20]. This effect is not accounted for by isospin breaking [21–24], but the accuracy of its calculation may be overestimated [25,26].

This problem seems to be solved after a recent analysis in [27] which showed that after correcting the  $\tau$  data for the missing  $\rho - \gamma$  mixing contribution, besides the other known isospin symmetry violating corrections, the  $\pi\pi$  I=1 part of the hadronic vacuum polarization contribution to the muon  $g - 2$  is fully compatible between  $\tau$  based and  $e^+e^-$  based evaluations including more recent BaBar [28] and KLOE [29] data. Further proof of the consistency of the data on  $\tau$  decays to two pions and  $e^+e^-$  annihilation is given by the global fit of the whole set of the  $\rho$ ,  $\omega$ , and  $\phi$  decays, taking into account mixing effects in the hidden local symmetry model [30].

# Meson Particle Listings

## $\rho(770)$

### References

1. J. Pisut and M. Roos, Nucl. Phys. **B6**, 325 (1968).
2. G.D. Lafferty, Z. Phys. **C60**, 659 (1993).
3. A. Abele *et al.*, Phys. Lett. **B402**, 195 (1997).
4. M. Benayoun *et al.*, Eur. Phys. J. **C31**, 525 (2003).
5. R. Barate *et al.*, Z. Phys. **C76**, 15 (1997).
6. L.M. Barkov *et al.*, Nucl. Phys. **B256**, 365 (1985).
7. S. Anderson *et al.*, Phys. Rev. **D61**, 112002 (2000).
8. M. Fujikawa *et al.*, Phys. Rev. **D78**, 072006 (2008).
9. S. Eidelman and V. Ivanchenko, Nucl. Phys. (Proc. Supp.) **B76**, 319 (1999).
10. K. Ackerstaff *et al.*, Eur. Phys. J. **C7**, 571 (1999).
11. M. Davier *et al.*, Nucl. Phys. (Proc. Supp.) **B123**, 47 (2003).
12. S. Schael *et al.*, Phys. Reports **421**, 191 (2005).
13. R.R. Akhmetshin *et al.*, Phys. Lett. **B527**, 161 (2002).
14. R.R. Akhmetshin *et al.*, Phys. Lett. **B578**, 285 (2004).
15. M. Davier *et al.*, Eur. Phys. J. **C27**, 497 (2003).
16. M. Davier *et al.*, Eur. Phys. J. **C31**, 503 (2003).
17. A. Aloisio *et al.*, Phys. Lett. **B606**, 12 (2005).
18. F. Ambrosino *et al.*, Phys. Lett. **B670**, 285 (2009).
19. M.N. Achasov *et al.*, Sov. Phys. JETP **101**, 1053 (2005).
20. M.N. Achasov *et al.*, Sov. Phys. JETP **103**, 380 (2006).
21. R. Alemany *et al.*, Eur. Phys. J. **C2**, 123 (1998).
22. H. Czyz and J.J. Kuhn, Eur. Phys. J. **C18**, 497 (2001).
23. V. Cirigliano *et al.*, Phys. Lett. **B513**, 361 (2001).
24. V. Cirigliano *et al.*, Eur. Phys. J. **C23**, 121 (2002).
25. K. Maltman and C.E. Wolfe, Phys. Rev. **D73**, 013004 (2006).
26. C.E. Wolfe and K. Maltman, Phys. Rev. **D80**, 114024 (2009).
27. F. Jegerlehner and R. Szafron, Eur. Phys. J. **C71**, 1632 (2011).
28. B. Aubert *et al.*, Phys. Rev. Lett. **103**, 231801 (2009).
29. F. Ambrosino *et al.*, Phys. Lett. **B700**, 102 (2011).
30. M. Benayoun *et al.*, Eur. Phys. J. **C72**, 1848 (2012).

### $\rho(770)$ MASS

We no longer list S-wave Breit-Wigner fits, or data with high combinatorial background.

### NEUTRAL ONLY, $e^+e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>775.26±0.25 OUR AVERAGE</b>				
775.02±0.35		1 LEES	12G	BABR $e^+e^- \rightarrow \pi^+\pi^-\gamma$
775.97±0.46±0.70	900k	2 AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
774.6±0.4±0.5	800k	3,4 ACHASOV	06	SND $e^+e^- \rightarrow \pi^+\pi^-$
775.65±0.64±0.50	114k	5,6 AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
775.9±0.5±0.5	1.98M	7 ALOISIO	03	KLOE $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.8±0.9±2.0	500k	7 ACHASOV	02	SND $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.9±1.1		8 BARKOV	85	OLYA $e^+e^- \rightarrow \pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
775.8±0.5±0.3	1.98M	9 ALOISIO	03	KLOE $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.9±0.6±0.5	1.98M	10 ALOISIO	03	KLOE $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.0±0.6±1.1	500k	11 ACHASOV	02	SND $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.1±0.7±5.3		12 BENAYOUN	98	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
770.5±1.9±5.1		13 GARDNER	98	RVUE $0.28-0.92 e^+e^- \rightarrow \pi^+\pi^-$
764.1±0.7		14 O'CONNELL	97	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
757.5±1.5		15 BERNICHA	94	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
768±1		16 GESKIN...	89	RVUE $e^+e^- \rightarrow \pi^+\pi^-$

### CHARGED ONLY, $\tau$ DECAYS and $e^+e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>775.11±0.34 OUR AVERAGE</b>					
774.6±0.2±0.5	5.4M	17,18 FUJIKAWA	08	BELL	$\pm$ $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
775.5±0.7		18,19 SCHAELE	05c	ALEP	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
775.5±0.5±0.4	1.98M	7 ALOISIO	03	KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.1±1.1±0.5	87k	20,21 ANDERSON	00A	CLE2	$\tau^- \rightarrow \pi^-\pi^-\pi^0\nu_\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
774.8±0.6±0.4	1.98M	10 ALOISIO	03	KLOE	$-$ $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
776.3±0.6±0.7	1.98M	10 ALOISIO	03	KLOE	$+$ $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
773.9±2.0±0.3		22 SANZ-CILLERCO	03	RVUE	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
774.5±0.7±1.5	500k	7 ACHASOV	02	SND	$\pm$ $1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.1±0.5		23 PICH	01	RVUE	$\tau^- \rightarrow \pi^-\pi^-\pi^0\nu_\tau$

### MIXED CHARGES, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>763.0±0.3±1.2</b>	600k	24 ABELE	99E	CBAR	$0 \pm 0.0 \bar{p}p \rightarrow \pi^+\pi^-\pi^0$

### CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>766.5±1.1 OUR AVERAGE</b>					
763.7±3.2		ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
768±9		AGUILAR...	91	EHS	400 $pp$
767±3	2935	25 CAPRARO	87	SPEC	$-$ 200 $\pi^- \text{Cu} \rightarrow \pi^-\pi^0\text{Cu}$
761±5	967	25 CAPRARO	87	SPEC	$-$ 200 $\pi^- \text{Pb} \rightarrow \pi^-\pi^0\text{Pb}$
771±4		HUSTON	86	SPEC	$+$ 202 $\pi^+\text{A} \rightarrow \pi^+\pi^0\text{A}$
766±7	6500	26 BYERLY	73	OSPK	$-$ 5 $\pi^-p$
766.8±1.5	9650	27 PISUT	68	RVUE	$-$ 1.7-3.2 $\pi^-p$ , $t < 10$
767±6	900	25 EISNER	67	HBC	$-$ 4.2 $\pi^-p$ , $t < 10$

### NEUTRAL ONLY, PHOTOPRODUCED

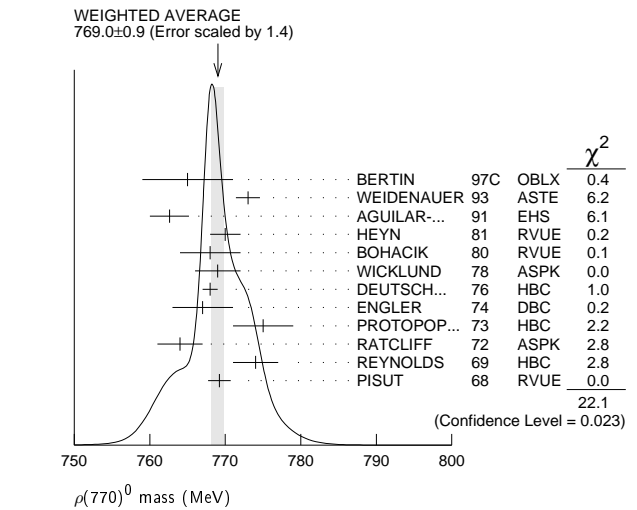
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>769.0±1.0 OUR AVERAGE</b>				
771±2±2	63.5k	28 ABRAMOWICZ12	ZEUS	$ep \rightarrow e\pi^+\pi^-p$
770±2±1	79k	29 BREITWEG	98B	ZEUS 50-100 $\gamma p$
767.6±2.7		BARTALUCCI	78	CNTR $\gamma p \rightarrow e^+\pi^-p$
775±5		GLADDING	73	CNTR 2.9-4.7 $\gamma p$
767±4	1930	BALLAM	72	HBC 2.8 $\gamma p$
770±4	2430	BALLAM	72	HBC 4.7 $\gamma p$
765±10		ALVENSLEB...	70	CNTR $\gamma\text{A}$ , $t < 0.01$
767.7±1.9	140k	BIGGS	70	CNTR $<4.1 \gamma\text{C} \rightarrow \pi^+\pi^-\text{C}$
765±5	4000	ASBURY	67B	CNTR $\gamma + \text{Pb}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
771±2	79k	30 BREITWEG	98B	ZEUS 50-100 $\gamma p$

### NEUTRAL ONLY, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>769.0±0.9 OUR AVERAGE</b>					Error includes scale factor of 1.4. See the ideogram below.
765±6		BERTIN	97c	OBLX	$0.0 \bar{p}p \rightarrow \pi^+\pi^-\pi^0$
773±1.6		WEIDENAUER	93	ASTE	$\bar{p}p \rightarrow \pi^+\pi^-\omega$
762.6±2.6		AGUILAR...	91	EHS	400 $pp$
770±2		31 HEYN	81	RVUE	Pion form factor
768±4		32,33 BOHACIK	80	RVUE	0
769±3		26 WICKLUND	78	ASPK	0 3,4,6 $\pi^\pm N$
768±1	76000	DEUTSCH...	76	HBC	0 16 $\pi^+p$
767±4	4100	ENGLER	74	DBC	0 6 $\pi^+n \rightarrow \pi^+\pi^-p$
775±4	32000	32 PROTOPOP...	73	HBC	0 7.1 $\pi^+p$ , $t < 0.4$
764±3	6800	RATCLIFF	72	ASPK	0 15 $\pi^-p$ , $t < 0.3$
774±3	1700	REYNOLDS	69	HBC	0 2.26 $\pi^-p$
769.2±1.5	13300	34 PISUT	68	RVUE	0 1.7-3.2 $\pi^-p$ , $t < 10$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
773.5±2.5		35 COLANGELO	01	RVUE	$\pi\pi \rightarrow \pi\pi$
762.3±0.5±1.2	600k	36 ABELE	99E	CBAR	0 $0.0 \bar{p}p \rightarrow \pi^+\pi^-\pi^0$
777±2	4943	37 ADAMS	97	E665	$470 \mu p \rightarrow \mu XB$
770±2		38 BOGOLYUB...	97	MIRA	$32 \bar{p}p \rightarrow \pi^+\pi^-\text{X}$
768±8		38 BOGOLYUB...	97	MIRA	$32 pp \rightarrow \pi^+\pi^-\text{X}$
761.1±2.9		DUBNICKA	89	RVUE	$\pi$ form factor
777.4±2.0		39 CHABAUD	83	ASPK	0 17 $\pi^-p$ polarized
769.5±0.7		32,33 LANG	79	RVUE	0
770±9		33 ESTABROOKS	74	RVUE	0 17 $\pi^-p \rightarrow \pi^+\pi^-\pi^0$
773.5±1.7	11200	25 JACOBS	72	HBC	0 2.8 $\pi^-p$
775±3	2250	HYAMS	68	OSPK	0 11.2 $\pi^-p$

# Meson Particle Listings

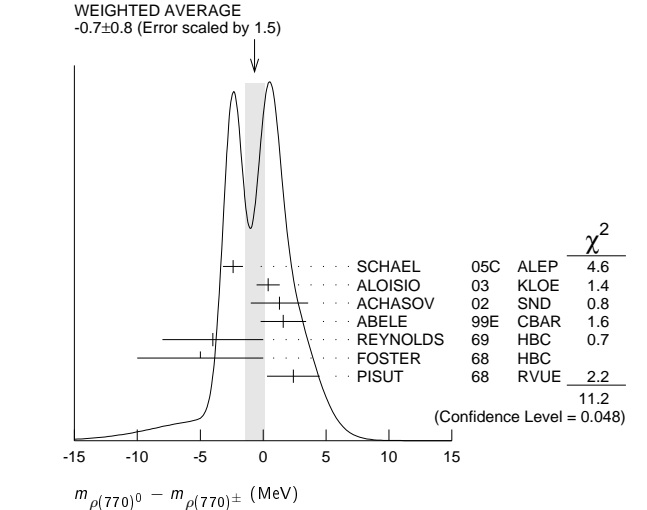
## $\rho(770)$



- <sup>1</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference and leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.
- <sup>2</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.
- <sup>3</sup> Supersedes ACHASOV 05A.
- <sup>4</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.
- <sup>5</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference.
- <sup>6</sup> Update of AKHMETSHIN 02.
- <sup>7</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .
- <sup>8</sup> From the GOUNARIS 68 parametrization of the pion form factor.
- <sup>9</sup> Assuming  $m_{\rho^+} = m_{\rho^-} = m_{\rho^0}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-} = \Gamma_{\rho^0}$ .
- <sup>10</sup> Without limitations on masses and widths.
- <sup>11</sup> Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0\pi\pi} = g_{\rho^\pm\pi\pi}$ .
- <sup>12</sup> Using the data of BARKOV 85 in the hidden local symmetry model.
- <sup>13</sup> From the fit to  $e^+e^- \rightarrow \pi^+\pi^-$  data from the compilations of HEYN 81 and BARKOV 85, including the GOUNARIS 68 parametrization of the pion form factor.
- <sup>14</sup> A fit of BARKOV 85 data assuming the direct  $\omega\pi\pi$  coupling.
- <sup>15</sup> Applying the S-matrix formalism to the BARKOV 85 data.
- <sup>16</sup> Includes BARKOV 85 data. Model-dependent width definition.
- <sup>17</sup>  $|F_\pi(0)|^2$  fixed to 1.
- <sup>18</sup> From the GOUNARIS 68 parametrization of the pion form factor.
- <sup>19</sup> The error combines statistical and systematic uncertainties. Supersedes BARATE 97M.
- <sup>20</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.
- <sup>21</sup> From the GOUNARIS 68 parametrization of the pion form factor. The second error is a model error taking into account different parametrizations of the pion form factor.
- <sup>22</sup> Using the data of BARATE 97M and the effective chiral Lagrangian.
- <sup>23</sup> From a fit of the model-independent parametrization of the pion form factor to the data of BARATE 97M.
- <sup>24</sup> Assuming the equality of  $\rho^+$  and  $\rho^-$  masses and widths.
- <sup>25</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.
- <sup>26</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.
- <sup>27</sup> From fit of 3-parameter relativistic P-wave Breit-Wigner to total mass distribution. Includes BATON 68, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65 and CARMONY 64.
- <sup>28</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.
- <sup>29</sup> From the parametrization according to SOEDING 66.
- <sup>30</sup> From the parametrization according to ROSS 66.
- <sup>31</sup> HEYN 81 includes all spacelike and timelike  $F_\pi$  values until 1978.
- <sup>32</sup> From pole extrapolation.
- <sup>33</sup> From phase shift analysis of GRAYER 74 data.
- <sup>34</sup> Includes MALAMUD 69, ARMENISE 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, GOLDHABER 64, ABOLINS 63.
- <sup>35</sup> Breit-Wigner mass from a phase-shift analysis of HYAMS 73 and PROTOPOPESCU 73 data.
- <sup>36</sup> Using relativistic Breit-Wigner and taking into account  $\rho$ - $\omega$  interference.
- <sup>37</sup> Systematic errors not evaluated.
- <sup>38</sup> Systematic effects not studied.
- <sup>39</sup> From fit of 3-parameter relativistic Breit-Wigner to helicity-zero part of P-wave intensity. CHABAUD 83 includes data of GRAYER 74.

$m_{\rho(770)^0} - m_{\rho(770)^\pm}$					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.7±0.8 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.			
-2.4±0.8		<sup>40</sup> SCHAE	05C	ALEP	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
0.4±0.7±0.6	1.98M	<sup>41</sup> ALOISIO	03	KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1.3±1.1±2.0	500k	<sup>41</sup> ACHASOV	02	SND	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1.6±0.6±1.7	600k	ABELE	99E	CBAR	$0.0 \overline{p}p \rightarrow \pi^+\pi^-\pi^0$
-4 ±4	3000	<sup>42</sup> REYNOLDS	69	HBC	$2.26 \pi^-p$
-5 ±5	3600	<sup>42</sup> FOSTER	68	HBC	$0.0 \overline{p}p$
2.4±2.1	22950	<sup>43</sup> PISUT	68	RVUE	$\pi N \rightarrow \rho N$

- <sup>40</sup> From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAE 05C and  $e^+e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05. Supersedes BARATE 97M.
- <sup>41</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .
- <sup>42</sup> From quoted masses of charged and neutral modes.
- <sup>43</sup> Includes MALAMUD 69, ARMENISE 68, BATON 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65, CARMONY 64, GOLDHABER 64, ABOLINS 63.



$m_{\rho(770)^+} - m_{\rho(770)^-}$					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.5±0.8±0.7	1.98M	<sup>44</sup> ALOISIO	03	KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$

<sup>44</sup> Without limitations on masses and widths.

$\rho(770)$ RANGE PARAMETER					
The range parameter $R$ enters an energy-dependent correction to the width, of the form $(1 + q^2 R^2) / (1 + q^2 R^2)$ , where $q$ is the momentum of one of the pions in the $\pi\pi$ rest system. At resonance, $q = q_r$ .					
VALUE (GeV <sup>-1</sup> )	DOCUMENT ID	TECN	CHG	COMMENT	
<b>5.3±0.9 -0.7</b>	CHABAUD	83	ASP K	0	17 $\pi^-p$ polarized

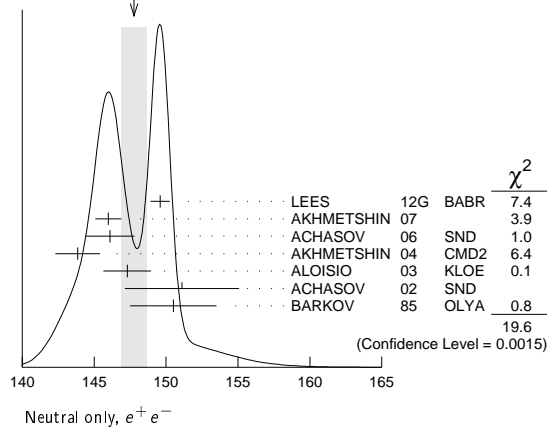
$\rho(770)$ WIDTH					
We no longer list S-wave Breit-Wigner fits, or data with high combinatorial background.					

NEUTRAL ONLY, $e^+e^-$					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>147.8 ±0.9 OUR AVERAGE</b>		Error includes scale factor of 2.0. See the ideogram below.			
149.59±0.67		<sup>45</sup> LEES	12G	BABR	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
145.98±0.75±0.50	900k	<sup>46</sup> AKHMETSHIN	07		$e^+e^- \rightarrow \pi^+\pi^-$
146.1 ±0.8 ±1.5	800k	<sup>47,48</sup> ACHASOV	06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
143.85±1.33±0.80	114k	<sup>49,50</sup> AKHMETSHIN	04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
147.3 ±1.5 ±0.7	1.98M	<sup>51</sup> ALOISIO	03	KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
151.1 ±2.6 ±3.0	500k	<sup>51</sup> ACHASOV	02	SND	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
150.5 ±3.0		<sup>52</sup> BARKOV	85	OLYA	$0 e^+e^- \rightarrow \pi^+\pi^-$
143.9 ±1.3 ±1.1	1.98M	<sup>53</sup> ALOISIO	03	KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
147.4 ±1.5 ±0.7	1.98M	<sup>54</sup> ALOISIO	03	KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
149.8 ±2.2 ±2.0	500k	<sup>55</sup> ACHASOV	02	SND	$1.02 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
147.9 ±1.5 ±7.5		<sup>56</sup> BENAYOUN	98	RVUE	$e^+e^- \rightarrow \pi^+\pi^-, \mu^+\mu^-$
153.5 ±1.3 ±4.6		<sup>57</sup> GARDNER	98	RVUE	$0.28-0.92 e^+e^- \rightarrow \pi^+\pi^-$
145.0 ±1.7		<sup>58</sup> O'CONNELL	97	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
142.5 ±3.5		<sup>59</sup> BERNICHA	94	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
138 ±1		<sup>60</sup> GESHKEN...	89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$



See key on page 601

## Meson Particle Listings

 $\rho(770)$ WEIGHTED AVERAGE  
147.8±0.9 (Error scaled by 2.0)CHARGED ONLY,  $\tau$  DECAYS and  $e^+e^-$ 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>149.1±0.8 OUR FIT</b>					
<b>149.1±0.8 OUR AVERAGE</b>					
148.1±0.4±1.7	5.4M	61,62 FUJIKAWA	08	BELL	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
149.0±1.2		62,63 SCHAEEL	05c	ALEP	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
149.9±2.3±2.0	500k	51 ACHASOV	02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
150.4±1.4±1.4	87k	64,65 ANDERSON	00A	CLE2	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
143.7±1.3±1.2	1.98M	51 ALOISIO	03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
142.9±1.3±1.4	1.98M	54 ALOISIO	03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
144.7±1.4±1.2	1.98M	54 ALOISIO	03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
150.2±2.0±0.7		66 SANZ-CILLERO	003	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
150.9±2.2±2.0	500k	55 ACHASOV	02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

## MIXED CHARGES, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>149.5±1.3</b>	600k	67 ABELE	99E	CBAR	$0 \pm 0.0 \bar{p} p \rightarrow \pi^+ \pi^- \pi^0$

## CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>150.2± 2.4 OUR FIT</b>					
<b>150.2± 2.4 OUR AVERAGE</b>					
152.8± 4.3		ABELE	97	CBAR	$\bar{p} n \rightarrow \pi^- \pi^0 \pi^0$
155 ± 11	2935	68 CAPRARO	87	SPEC	$200 \pi^- \text{Cu} \rightarrow \pi^- \pi^0 \text{Cu}$
154 ± 20	967	68 CAPRARO	87	SPEC	$200 \pi^- \text{Pb} \rightarrow \pi^- \pi^0 \text{Pb}$
150 ± 5		HUSTON	86	SPEC	$202 \pi^+ \text{A} \rightarrow \pi^+ \pi^0 \text{A}$
146 ± 12	6500	69 BYERLY	73	OSPK	$5 \pi^- p$
148.2± 4.1	965.0	70 PISUT	68	RVUE	$1.7-3.2 \pi^- p, t < 10$
146 ± 13	900	EISNER	67	HBC	$4.2 \pi^- p, t < 10$

## NEUTRAL ONLY, PHOTOPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>151.7± 2.6 OUR AVERAGE</b>				
155 ± 5 ± 2	63.5k	71 ABRA MOWICZ	12	ZEUS $e p \rightarrow e^+ \pi^- \pi^0$
146 ± 3 ± 13	79k	72 BREITWEG	98B	ZEUS $50-100 \gamma p$
150.9± 3.0		BARTALUCCI	78	CNTR $\gamma p \rightarrow e^+ e^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
138 ± 3	79k	73 BREITWEG	98B	ZEUS $50-100 \gamma p$
147 ± 11		GLADDING	73	CNTR $2.9-4.7 \gamma p$
155 ± 12	2430	BALLAM	72	HBC $4.7 \gamma p$
145 ± 13	1930	BALLAM	72	HBC $2.8 \gamma p$
140 ± 5		ALVENSLEB...	70	CNTR $\gamma A, t < 0.01$
146.1± 2.9	140k	BIGGS	70	CNTR $< 4.1 \gamma C \rightarrow \pi^+ \pi^- C$
160 ± 10		LANZEROTTI	68	CNTR $\gamma p$
130 ± 5	4000	ASBURY	67B	CNTR $\gamma + \text{Pb}$

## NEUTRAL ONLY, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>150.9± 1.7 OUR AVERAGE</b>					
Error includes scale factor of 1.1.					
122 ± 20		BERTIN	97c	OBLX	$0.0 \bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
145.7± 5.3		WEIDENAUER	93	ASTE	$\bar{p} p \rightarrow \pi^+ \pi^- \omega$
144.9± 3.7		DUBNICKA	89	RVUE	$\pi$ form factor
148 ± 6	74,75	BOHACIK	80	RVUE	0
152 ± 9		WICKLUND	78	ASPK	$3.4, 6 \pi^\pm p N$
154 ± 2	76000	DEUTSCH...	76	HBC	$16 \pi^+ p$
157 ± 8	6800	RATCLIFF	72	ASPK	$15 \pi^- p, t < 0.3$
143 ± 8	1700	REYNOLDS	69	HBC	$2.26 \pi^- p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

147.0± 2.5	600k	76 ABELE	99E	CBAR	$0.0 \bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
146 ± 3	4943	77 ADAMS	97	E665	$470 \mu p \rightarrow \mu X B$
160.0± 4.1		78 CHABAUD	83	ASPK	$17 \pi^- p$ polarized
155 ± 1		79 HEYN	81	RVUE	$\pi$ form factor
148.0± 1.3		74,75 LANG	79	RVUE	0
146 ± 14	4100	ENGLER	74	DBC	$6 \pi^+ n \rightarrow \pi^+ \pi^- p$
143 ± 13		75 ESTABROOKS	74	RVUE	$17 \pi^- p \rightarrow \pi^+ \pi^- n$
160 ± 10	32000	74 PROTOPOP...	73	HBC	$7.1 \pi^+ p, t < 0.4$
145 ± 12	2250	68 HYAMS	68	OSPK	$11.2 \pi^- p$
163 ± 15	13300	80 PISUT	68	RVUE	$1.7-3.2 \pi^- p, t < 10$

45 Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference and leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

46 A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.

47 Supersedes ACHASOV 05A.

48 A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.

49 Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference.

50 From a fit in the energy range 0.61 to 0.96 GeV. Update of AKHMETSHIN 02.

51 Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

52 From the GOUNARIS 68 parametrization of the pion form factor.

53 Assuming  $m_{\rho^+} = m_{\rho^-} = m_{\rho^0}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-} = \Gamma_{\rho^0}$ .

54 Without limitations on masses and widths.

55 Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0 \pi \pi} = g_{\rho^\pm \pi \pi}$ .

56 Using the data of BARKOV 85 in the hidden local symmetry model.

57 From the fit to  $e^+ e^- \rightarrow \pi^+ \pi^-$  data from the compilations of HEYN 81 and BARKOV 85, including the GOUNARIS 68 parametrization of the pion form factor.

58 A fit of BARKOV 85 data assuming the direct  $\omega \pi \pi$  coupling.

59 Applying the S-matrix formalism to the BARKOV 85 data.

60 Includes BARKOV 85 data. Model-dependent width definition.

61  $|F_\pi(0)|^2$  fixed to 1.

62 From the GOUNARIS 68 parametrization of the pion form factor.

63 The error combines statistical and systematic uncertainties. Supersedes BARATE 97M.

64  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.

65 From the GOUNARIS 68 parametrization of the pion form factor. The second error is a model error taking into account different parametrizations of the pion form factor.

66 Using the data of BARATE 97M and the effective chiral Lagrangian.

67 Assuming the equality of  $\rho^+$  and  $\rho^-$  masses and widths.

68 Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

69 Phase shift analysis. Systematic errors added corresponding to spread of different fits.

70 From fit of 3-parameter relativistic  $P$ -wave Breit-Wigner to total mass distribution. Includes BATON 68, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65 and CARMONY 64.

71 Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.

72 From the parametrization according to SOEDING 66.

73 From the parametrization according to ROSS 66.

74 From pole extrapolation.

75 From phase shift analysis of GRAYER 74 data.

76 Using relativistic Breit-Wigner and taking into account  $\rho$ - $\omega$  interference.

77 Systematic errors not evaluated.

78 From fit of 3-parameter relativistic Breit-Wigner to helicity-zero part of  $P$ -wave intensity. CHABAUD 83 includes data of GRAYER 74.

79 HEYN 81 includes all spacelike and timelike  $F_\pi$  values until 1978.

80 Includes MALAMUD 69, ARMENISE 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, GOLDBABER 64, ABOLINS 63.

 $\Gamma_{\rho(770)^0} - \Gamma_{\rho(770)^\pm}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3±1.3 OUR AVERAGE</b>				
Error includes scale factor of 1.4.				
-0.2±1.0		81 SCHAEEL	05c	ALEP $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
3.6±1.8±1.7	1.98M	82 ALOISIO	03	KLOE $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

 $\Gamma_{\rho(770)^+} - \Gamma_{\rho(770)^-}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.8±2.0±0.5</b>				
	1.98M	83 ALOISIO	03	KLOE $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

81 From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEEL 05c and  $e^+ e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05. Supersedes BARATE 97M.

82 Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

83 Without limitations on masses and widths.

 $\rho(770)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1 \quad \pi \pi$	$\sim 100$	%
<b><math>\rho(770)^\pm</math> decays</b>		
$\Gamma_2 \quad \pi^\pm \pi^0$	$\sim 100$	%
$\Gamma_3 \quad \pi^\pm \gamma$	( 4.5 ± 0.5 )	$\times 10^{-4}$ S=2.2
$\Gamma_4 \quad \pi^\pm \eta$	< 6	$\times 10^{-3}$ CL=84%
$\Gamma_5 \quad \pi^\pm \pi^+ \pi^- \pi^0$	< 2.0	$\times 10^{-3}$ CL=84%

## Meson Particle Listings

 $\rho(770)$  $\rho(770)^0$  decays

	$\sim 100$	%
$\Gamma_6 \pi^+ \pi^-$	( 9.9 $\pm$ 1.6 )	$\times 10^{-3}$
$\Gamma_7 \pi^+ \pi^- \gamma$	( 6.0 $\pm$ 0.8 )	$\times 10^{-4}$
$\Gamma_8 \pi^0 \gamma$	( 3.00 $\pm$ 0.20 )	$\times 10^{-4}$
$\Gamma_9 \eta \gamma$	( 4.5 $\pm$ 0.8 )	$\times 10^{-5}$
$\Gamma_{10} \pi^0 \pi^0 \gamma$	( 4.55 $\pm$ 0.28 )	$\times 10^{-5}$
$\Gamma_{11} \mu^+ \mu^-$	[a] ( 4.72 $\pm$ 0.05 )	$\times 10^{-5}$
$\Gamma_{12} e^+ e^-$	( 1.01 $\pm$ 0.54 $\pm$ 0.34 )	$\times 10^{-4}$
$\Gamma_{13} \pi^+ \pi^- \pi^0$	( 1.8 $\pm$ 0.9 )	$\times 10^{-5}$
$\Gamma_{14} \pi^+ \pi^- \pi^+ \pi^-$	( 1.6 $\pm$ 0.8 )	$\times 10^{-5}$
$\Gamma_{15} \pi^+ \pi^- \pi^0 \pi^0$	< 1.2	$\times 10^{-5}$
$\Gamma_{16} \pi^0 e^+ e^-$		
$\Gamma_{17} \eta e^+ e^-$		

[a] The  $\omega\rho$  interference is then due to  $\omega\rho$  mixing only, and is expected to be small. If  $e\mu$  universality holds,  $\Gamma(\rho^0 \rightarrow \mu^+ \mu^-) = \Gamma(\rho^0 \rightarrow e^+ e^-) \times 0.99785$ .

## CONSTRAINED FIT INFORMATION

An overall fit to the total width and a partial width uses 10 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 10.7$  for 8 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	-100
$\Gamma$	15 -15
	$x_2 \quad x_3$

Mode	Rate (MeV)	Scale factor
$\Gamma_2 \pi^\pm \pi^0$	150.2 $\pm$ 2.4	
$\Gamma_3 \pi^\pm \gamma$	0.068 $\pm$ 0.007	2.3

## CONSTRAINED FIT INFORMATION

An overall fit to the total width, a partial width, and 7 branching ratios uses 21 measurements and one constraint to determine 9 parameters. The overall fit has a  $\chi^2 = 6.0$  for 13 degrees of freedom.

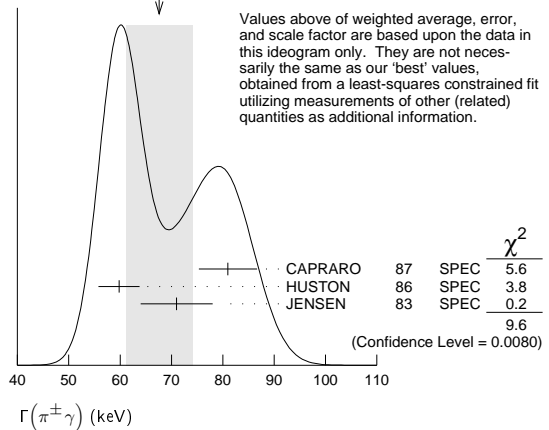
The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_7$	-100						
$x_8$	-5	0					
$x_9$	-1	0	1				
$x_{10}$	-1	0	0	0			
$x_{11}$	2	-3	0	0	0		
$x_{12}$	0	0	-8	-9	0	0	
$x_{14}$	-1	0	0	0	0	0	0
$\Gamma$	0	0	4	5	0	0	-54
	$x_6$	$x_7$	$x_8$	$x_9$	$x_{10}$	$x_{11}$	$x_{12}$

Mode	Rate (MeV)
$\Gamma_6 \pi^+ \pi^-$	147.5 $\pm$ 0.9
$\Gamma_7 \pi^+ \pi^- \gamma$	1.48 $\pm$ 0.24
$\Gamma_8 \pi^0 \gamma$	0.089 $\pm$ 0.012
$\Gamma_9 \eta \gamma$	0.0447 $\pm$ 0.0031
$\Gamma_{10} \pi^0 \pi^0 \gamma$	0.0066 $\pm$ 0.0012
$\Gamma_{11} \mu^+ \mu^-$	[a] 0.0068 $\pm$ 0.0004
$\Gamma_{12} e^+ e^-$	[a] 0.00704 $\pm$ 0.00006
$\Gamma_{14} \pi^+ \pi^- \pi^+ \pi^-$	0.0027 $\pm$ 0.0014

 $\rho(770)$  PARTIAL WIDTHS

$\Gamma(\pi^\pm \gamma)$	VALUE (keV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>68 <math>\pm</math> 7 OUR FIT</b>		Error includes scale factor of 2.3.			
<b>68 <math>\pm</math> 7 OUR AVERAGE</b>		Error includes scale factor of 2.2. See the ideogram below.			
81 $\pm$ 4 $\pm$ 4	CAPRARO	87	SPEC	-	200 $\pi^- A \rightarrow \pi^- \pi^0 A$
59.8 $\pm$ 4.0	HUSTON	86	SPEC	+	202 $\pi^+ A \rightarrow \pi^+ \pi^0 A$
71 $\pm$ 7	JENSEN	83	SPEC	-	156-260 $\pi^- A \rightarrow \pi^- \pi^0 A$

WEIGHTED AVERAGE  
68 $\pm$ 7 (Error scaled by 2.2) $\Gamma(e^+ e^-)$ 

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.04 <math>\pm</math> 0.06 OUR FIT</b>				
<b>7.04 <math>\pm</math> 0.06 OUR AVERAGE</b>				
7.048 $\pm$ 0.057 $\pm$ 0.050	900k	84 AKHMETSHIN 07		$e^+ e^- \rightarrow \pi^+ \pi^-$
7.06 $\pm$ 0.11 $\pm$ 0.05	114k	85,86 AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^-$
6.77 $\pm$ 0.10 $\pm$ 0.30		BARKOV 85	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7.12 $\pm$ 0.02 $\pm$ 0.11	800k	87 ACHASOV 06	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
6.3 $\pm$ 0.1		88 BENAYOUN 98	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$ , $\mu^+ \mu^-$

 $\Gamma(\pi^0 \gamma)$ 

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
77 $\pm$ 17 $\pm$ 11	36500	89 ACHASOV 03	SND	0.60-0.97 $e^+ e^- \rightarrow \pi^0 \gamma$
121 $\pm$ 31		DOLINSKY 89	ND	$e^+ e^- \rightarrow \pi^0 \gamma$

 $\Gamma(\eta \gamma)$ 

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
62 $\pm$ 17	90 DOLINSKY 89	ND	$e^+ e^- \rightarrow \eta \gamma$

 $\Gamma(\pi^+ \pi^- \pi^+ \pi^-)$ 

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.8 $\pm$ 1.4 $\pm$ 0.5	153	AKHMETSHIN 00	CMD2	0.6-0.97 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^-$
84 A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.				
85 Using the GOUNARIS 68 parametrization with the complex phase of the $\rho$ - $\omega$ interference.				
86 From a fit in the energy range 0.61 to 0.96 GeV. Update of AKHMETSHIN 02.				
87 Supersedes ACHASOV 05A.				
88 Using the data of BARKOV 85 in the hidden local symmetry model.				
89 Using $\Gamma_{\text{total}} = 147.9 \pm 1.3$ MeV and $B(\rho \rightarrow \pi^0 \gamma)$ from ACHASOV 03.				
90 Solution corresponding to constructive $\omega$ - $\rho$ interference.				

 $\rho(770) \Gamma(e^+ e^-) \Gamma(i) / \Gamma^2(\text{total})$ 

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma \times \Gamma_6/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.876 ± 0.023 ± 0.064</b>	800k	91,92 ACHASOV	06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
4.72 ± 0.02		93 BENAYOUN	10	RVUE	0.4-1.05 $e^+e^-$

91 Supersedes ACHASOV 05A.  
92 A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.  
93 A simultaneous fit of  $e^+ e^- \rightarrow \pi^+ \pi^-$ ,  $\pi^+ \pi^- \pi^0$ ,  $\pi^0 \gamma$ ,  $\eta \gamma$  data.

 $\Gamma(e^+ e^-) / \Gamma_{\text{total}} \times \Gamma(\eta \gamma) / \Gamma_{\text{total}}$ 

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.42 <math>\pm</math> 0.10 OUR FIT</b>				
<b>1.45 <math>\pm</math> 0.12 OUR AVERAGE</b>				
1.32 $\pm$ 0.14 $\pm$ 0.08	33k	94 ACHASOV 07B	SND	0.6-1.38 $e^+ e^- \rightarrow \eta \gamma$
1.50 $\pm$ 0.65 $\pm$ 0.09	17.4k	95 AKHMETSHIN 05	CMD2	0.60-1.38 $e^+ e^- \rightarrow \eta \gamma$
1.61 $\pm$ 0.20 $\pm$ 0.11	23k	96,97 AKHMETSHIN 01B	CMD2	$e^+ e^- \rightarrow \eta \gamma$
1.85 $\pm$ 0.49		98 DOLINSKY 89	ND	$e^+ e^- \rightarrow \eta \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.05 $\pm$ 0.02		99 BENAYOUN 10	RVUE	0.4-1.05 $e^+ e^-$



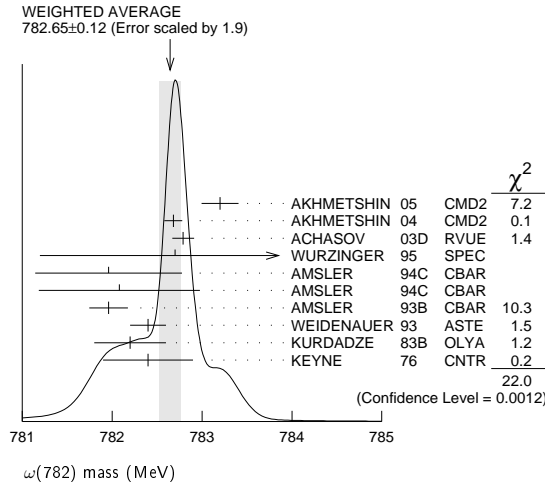


See key on page 601

# Meson Particle Listings

$\omega(782)$

- <sup>5</sup> Observed by threshold-crossing technique. Mass resolution = 4.8 MeV FWHM.  
<sup>6</sup> From the  $\rho-\omega$  interference in the  $\pi^+\pi^-$  mass spectrum using the Breit-Wigner for the  $\omega$  and leaving its mass and width as free parameters of the fit.  
<sup>7</sup> Systematic uncertainties underestimated.  
<sup>8</sup> From best-resolution sample of COYNE 71.  
<sup>9</sup> From  $\rho-\omega$  interference in the  $\pi^+\pi^-$  mass spectrum assuming  $\omega$  width 12.6 MeV.



## $\omega(782)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.49±0.08 OUR AVERAGE</b>				
8.68±0.23±0.10	11200	<sup>1</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
8.68±0.04±0.15	1.2M	<sup>2</sup> ACHASOV 03D	RVUE	0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
8.2 ±0.3	19500	WURZINGER 95	SPEC	1.33 $p d \rightarrow {}^3\text{He} \omega$
8.4 ±0.1		<sup>3</sup> AULCHENKO 87	ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
8.30±0.40		BARKOV 87	CMD	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
9.8 ±0.9	1488	KURDADZE 83B	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
9.0 ±0.8	433	CORDIER 80	DM1	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
9.1 ±0.8	451	BENAKSAS 72B	OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.13±0.45		<sup>4</sup> LEES 12G	BABR	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
12 ±2	1430	COOPER 78B	HBC	0.7–0.8 $\bar{p} p \rightarrow 5\pi$
9.4 ±2.5	2100	GESSAROLI 77	HBC	11 $\pi^- p \rightarrow \omega n$
10.22±0.43	20000	<sup>5</sup> KEYNE 76	CNTR	$\pi^- p \rightarrow \omega n$
13.3 ±2	418	AGUILAR-... 72B	HBC	3.9,4.6 $K^- p$
10.5 ±1.5		BORENSTEIN 72	HBC	2.18 $K^- p$
7.70±0.9 ±1.15	940	BROWN 72	MMS	2.5 $\pi^- p \rightarrow n\text{MM}$
10.3 ±1.4	510	BIZZARRI 71	HBC	0.0 $p\bar{p} \rightarrow K_1^0 K_1^0 \omega$
12.8 ±3.0	248	BIZZARRI 71	HBC	0.0 $p\bar{p} \rightarrow K^+ K^- \omega$
9.5 ±1.0	3583	COYNE 71	HBC	3.7 $\pi^+ p \rightarrow \rho \pi^+ \pi^+ \pi^- \pi^0$

- <sup>1</sup> Update of AKHMETSHIN 00c.  
<sup>2</sup> From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the  $\pi^+\pi^-\pi^0$  and ANTONELLI 92 on the  $\omega\pi^+\pi^-$  final states. Supersedes ACHASOV 99E and ACHASOV 02E.  
<sup>3</sup> Relativistic Breit-Wigner includes radiative corrections.  
<sup>4</sup> From the  $\rho-\omega$  interference in the  $\pi^+\pi^-$  mass spectrum using the Breit-Wigner for the  $\omega$  and leaving its mass and width as free parameters of the fit.  
<sup>5</sup> Observed by threshold-crossing technique. Mass resolution = 4.8 MeV FWHM.

## $\omega(782)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $\pi^+\pi^-\pi^0$	(89.2 ±0.7) %	
$\Gamma_2$ $\pi^0\gamma$	( 8.28±0.28) %	S=2.1
$\Gamma_3$ $\pi^+\pi^-$	( 1.53 <sup>+0.11</sup> <sub>-0.13</sub> ) %	S=1.2
$\Gamma_4$ neutrals (excluding $\pi^0\gamma$ )	( 8 <sup>+8</sup> <sub>-5</sub> ) ×10 <sup>-3</sup>	S=1.1
$\Gamma_5$ $\eta\gamma$	( 4.6 ±0.4 ) ×10 <sup>-4</sup>	S=1.1
$\Gamma_6$ $\pi^0 e^+ e^-$	( 7.7 ±0.6 ) ×10 <sup>-4</sup>	
$\Gamma_7$ $\pi^0 \mu^+ \mu^-$	( 1.3 ±0.4 ) ×10 <sup>-4</sup>	S=2.1
$\Gamma_8$ $\eta e^+ e^-$		
$\Gamma_9$ $e^+ e^-$	( 7.28±0.14) ×10 <sup>-5</sup>	S=1.3
$\Gamma_{10}$ $\pi^+\pi^-\pi^0\pi^0$	< 2 ×10 <sup>-4</sup>	CL=90%
$\Gamma_{11}$ $\pi^+\pi^-\gamma$	< 3.6 ×10 <sup>-3</sup>	CL=95%

$\Gamma_{12}$ $\pi^+\pi^-\pi^+\pi^-$	< 1 ×10 <sup>-3</sup>	CL=90%
$\Gamma_{13}$ $\pi^0\pi^0\gamma$	( 6.6 ±1.1 ) ×10 <sup>-5</sup>	
$\Gamma_{14}$ $\eta\pi^0\gamma$	< 3.3 ×10 <sup>-5</sup>	CL=90%
$\Gamma_{15}$ $\mu^+\mu^-$	( 9.0 ±3.1 ) ×10 <sup>-5</sup>	
$\Gamma_{16}$ $3\gamma$	< 1.9 ×10 <sup>-4</sup>	CL=95%

## Charge conjugation (C) violating modes

$\Gamma_{17}$ $\eta\pi^0$	C < 2.1 ×10 <sup>-4</sup>	CL=90%
$\Gamma_{18}$ $2\pi^0$	C < 2.1 ×10 <sup>-4</sup>	CL=90%
$\Gamma_{19}$ $3\pi^0$	C < 2.3 ×10 <sup>-4</sup>	CL=90%

## CONSTRAINED FIT INFORMATION

An overall fit to 15 branching ratios uses 51 measurements and one constraint to determine 10 parameters. The overall fit has a  $\chi^2 = 51.8$  for 42 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	22								
$x_3$	-18	-4							
$x_4$	-92	-56	1						
$x_5$	7	7	-1	-9					
$x_6$	-1	0	0	0	0				
$x_7$	-1	0	0	0	0	0			
$x_9$	-38	-33	7	44	-21	0	0		
$x_{13}$	1	4	0	-2	0	0	0	-1	
$x_{15}$	0	0	0	0	0	0	0	0	0
	$x_1$	$x_2$	$x_3$	$x_4$	$x_5$	$x_6$	$x_7$	$x_9$	$x_{13}$

## $\omega(782)$ PARTIAL WIDTHS

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
880±50	7815	<sup>1</sup> ACHASOV 13	SND	1.05–2.00 $e^+e^- \rightarrow \pi^0\pi^0\gamma$	
788±12±27	36500	<sup>2</sup> ACHASOV 03	SND	0.60–0.97 $e^+e^- \rightarrow \pi^0\gamma$	
764±51	10625	DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$	
<sup>1</sup> Systematic uncertainty not estimated. <sup>2</sup> Using $\Gamma_\omega = 8.44 \pm 0.09$ MeV and $B(\omega \rightarrow \pi^0\gamma)$ from ACHASOV 03.					

VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_5$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.1±2.5	<sup>1</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \eta\gamma$	
<sup>1</sup> Using $\Gamma_\omega = 8.4 \pm 0.1$ MeV and $B(\omega \rightarrow \eta\gamma)$ from DOLINSKY 89.				

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_9$
<b>0.60 ±0.02 OUR EVALUATION</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.591±0.015	11200	<sup>1,2</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
0.653±0.003±0.021	1.2M	<sup>3</sup> ACHASOV 03D	RVUE	0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
0.600±0.031	10625	DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$	
<sup>1</sup> Using $B(\omega \rightarrow \pi^+\pi^-\pi^0) = 0.891 \pm 0.007$ and $\Gamma_{\text{total}} = 8.44 \pm 0.09$ MeV. <sup>2</sup> Update of AKHMETSHIN 00c. <sup>3</sup> Using ACHASOV 03, ACHASOV 03D and $B(\omega \rightarrow \pi^+\pi^-) = (1.70 \pm 0.28)\%$ .					

## $\omega(782)$ $\Gamma(e^+e^-)\Gamma(i)/\Gamma^2(\text{total})$

VALUE (units 10 <sup>-5</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma \times \Gamma_1/\Gamma$
<b>6.49±0.11 OUR FIT</b>				Error includes scale factor of 1.3.	
<b>6.38±0.10 OUR AVERAGE</b>				Error includes scale factor of 1.1.	
6.24±0.11±0.08	11.2k	<sup>1</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
6.70±0.06±0.27		AUBERT,B 04N	BABR	10.6 $e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma$	
6.74±0.04±0.24	1.2M	<sup>2,3</sup> ACHASOV 03D	RVUE	0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
6.37±0.35		<sup>2</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
6.45±0.24		<sup>2</sup> BARKOV 87	CMD	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
5.79±0.42	1488	<sup>2</sup> KURDADZE 83B	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
5.89±0.54	433	<sup>2</sup> CORDIER 80	DM1	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
7.54±0.84	451	<sup>2</sup> BENAKSAS 72B	OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
6.20±0.13		<sup>4</sup> BENAYOUN 10	RVUE	0.4–1.05 $e^+e^-$	

# Meson Particle Listings

## $\omega(782)$

- <sup>1</sup> Update of AKHMETSHIN 00c.  
<sup>2</sup> Recalculated by us from the cross section in the peak.  
<sup>3</sup> From the combined fit of ANTONELLI 92, ACHASOV 01e, ACHASOV 02e, and ACHASOV 03d data on the  $\pi^+\pi^-\pi^0$  and ANTONELLI 92 on the  $\omega\pi^+\pi^-$  final states. Supersedes ACHASOV 99e and ACHASOV 02e.  
<sup>4</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \qquad \Gamma_9/\Gamma \times \Gamma_2/\Gamma$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.02±0.20 OUR FIT</b>	Error includes scale factor of 1.9.			
<b>6.45±0.17 OUR AVERAGE</b>				
6.47±0.14±0.39	18680	AKHMETSHIN 05	CMD2	0.60-1.38 $e^+e^- \rightarrow \pi^0\gamma$
6.50±0.11±0.20	36500	<sup>1</sup> ACHASOV 03	SND	0.60-0.97 $e^+e^- \rightarrow \pi^0\gamma$
6.34±0.21±0.21	10625	<sup>2</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.80±0.13		<sup>3</sup> BENAYOUN 10	RVUE	0.4-1.05 $e^+e^-$
<sup>1</sup> Using $\sigma_\phi \rightarrow \pi^0\gamma$ from ACHASOV 00 and $m_\omega = 782.57$ MeV in the model with the energy-independent phase of $\rho$ - $\omega$ interference equal to $(-10.2 \pm 7.0)^\circ$ .				
<sup>2</sup> Recalculated by us from the cross section in the peak.				
<sup>3</sup> A simultaneous fit of $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$ data.				

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+\pi^-)/\Gamma_{\text{total}} \qquad \Gamma_9/\Gamma \times \Gamma_3/\Gamma$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.225±0.058±0.041</b>	800k	<sup>1</sup> ACHASOV 06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.166±0.036		<sup>2</sup> BENAYOUN 13	RVUE	0.4-1.05 $e^+e^-$
1.05 ±0.08		<sup>3</sup> DAVIER 13	RVUE	$e^+e^- \rightarrow \pi^+\pi^-(\gamma)$
<sup>1</sup> Supersedes ACHASOV 05A.				
<sup>2</sup> A simultaneous fit to $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma, K\bar{K}$ , and $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$ data. Supersedes BENAYOUN 10.				
<sup>3</sup> From $e^+e^- \rightarrow \pi^+\pi^-(\gamma)$ data of LEES 12g.				

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\eta\gamma)/\Gamma_{\text{total}} \qquad \Gamma_9/\Gamma \times \Gamma_5/\Gamma$

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.32±0.28 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>3.18±0.28 OUR AVERAGE</b>				
3.10±0.31±0.11	33k	<sup>1</sup> ACHASOV 07B	SND	0.6-1.38 $e^+e^- \rightarrow \eta\gamma$
3.17 $^{+1.85}_{-1.31}$ ±0.21	17.4k	<sup>2</sup> AKHMETSHIN 05	CMD2	0.60-1.38 $e^+e^- \rightarrow \eta\gamma$
3.41±0.52±0.21	23k	<sup>3,4</sup> AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.50±0.10		<sup>5</sup> BENAYOUN 10	RVUE	0.4-1.05 $e^+e^-$
<sup>1</sup> From a combined fit of $\sigma(e^+e^- \rightarrow \eta\gamma)$ with $\eta \rightarrow 3\pi^0$ and $\eta \rightarrow \pi^+\pi^-\pi^0$ , and fixing $B(\eta \rightarrow 3\pi^0) / B(\eta \rightarrow \pi^+\pi^-\pi^0) = 1.44 \pm 0.04$ . Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.				
<sup>2</sup> From the $\eta \rightarrow 2\gamma$ decay and using $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .				
<sup>3</sup> From the $\eta \rightarrow 3\pi^0$ decay and using $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .				
<sup>4</sup> The combined fit from 600 to 1380 MeV taking into account $\rho(770), \omega(782), \phi(1020)$ , and $\rho(1450)$ (mass and width fixed at 1450 MeV and 310 MeV respectively).				
<sup>5</sup> A simultaneous fit of $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$ data.				

### $\omega(782)$ BRANCHING RATIOS

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}} \qquad \Gamma_1/\Gamma$

NIECKING 12 describes final-state interactions between the three pions in a dispersive framework using data on the  $\pi\pi P$ -wave scattering phase shift.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.9024±0.0019		<sup>1</sup> AMBROSINO 08G	KLOE	1.0-1.03 $e^+e^- \rightarrow \pi^+\pi^-\pi^0, 2\pi^0\gamma$
0.8965±0.0016±0.0048	1.2M	<sup>2,3</sup> ACHASOV 03D	RVUE	0.44-2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.880 ±0.020 ±0.032	11200	<sup>3,4</sup> AKHMETSHIN 00C	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.8942±0.0062		<sup>3</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<sup>1</sup> Not independent of $\Gamma(\pi^0\gamma) / \Gamma(\pi^+\pi^-\pi^0)$ from AMBROSINO 08G.				
<sup>2</sup> Using ACHASOV 03, ACHASOV 03D and $B(\omega \rightarrow \pi^+\pi^-) = (1.70 \pm 0.28)\%$ .				
<sup>3</sup> Not independent of the corresponding $\Gamma(e^+e^-) \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}^2$ .				
<sup>4</sup> Using $\Gamma(e^+e^-) = 0.60 \pm 0.02$ keV.				

$\Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \qquad \Gamma_2/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.09±0.14		<sup>1</sup> AMBROSINO 08G	KLOE	$e^+e^- \rightarrow \pi^+\pi^-\pi^0, 2\pi^0\gamma$
9.06±0.20±0.57	18680	<sup>2,3</sup> AKHMETSHIN 05	CMD2	0.60-1.38 $e^+e^- \rightarrow \pi^0\gamma$
9.34±0.15±0.31	36500	<sup>3</sup> ACHASOV 03	SND	0.60-0.97 $e^+e^- \rightarrow \pi^0\gamma$
8.65±0.16±0.42	1.2M	<sup>4,5</sup> ACHASOV 03D	RVUE	0.44-2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
8.39±0.24	9975	<sup>6</sup> BENAYOUN 96	RVUE	$e^+e^- \rightarrow \pi^0\gamma$
8.88±0.62	10625	<sup>3</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$

- <sup>1</sup> Not independent of  $\Gamma(\pi^0\gamma) / \Gamma(\pi^+\pi^-\pi^0)$  from AMBROSINO 08G.  
<sup>2</sup> Using  $B(\omega \rightarrow e^+e^-) = (7.14 \pm 0.13) \times 10^{-5}$ .  
<sup>3</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .  
<sup>4</sup> Using ACHASOV 03, ACHASOV 03D and  $B(\omega \rightarrow \pi^+\pi^-) = (1.70 \pm 0.28)\%$ .  
<sup>5</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}^2$ .  
<sup>6</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, DOLINSKY 91 taking into account the triangle anomaly contributions.

$\Gamma(\pi^0\gamma)/\Gamma(\pi^+\pi^-\pi^0) \qquad \Gamma_2/\Gamma_1$

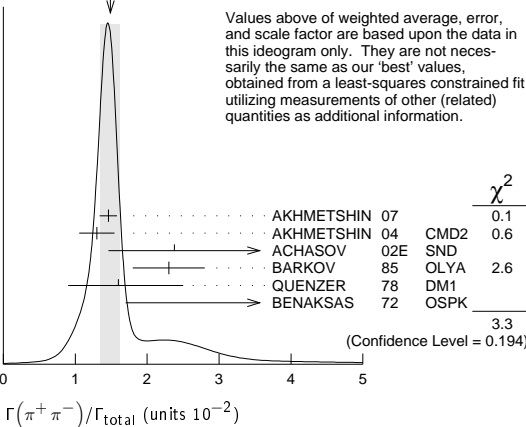
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>9.28±0.31 OUR FIT</b>	Error includes scale factor of 2.3.		
<b>9.05±0.27 OUR AVERAGE</b>	Error includes scale factor of 1.8.		
8.97±0.16	AMBROSINO 08G	KLOE	$e^+e^- \rightarrow \pi^+\pi^-\pi^0, 2\pi^0\gamma$
9.94±0.36±0.38	<sup>1</sup> AULCHENKO 00A	SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0, 2\pi^0\gamma$
8.4 ±1.3	KEYNE 76	CNTR	$\pi^-\rho \rightarrow \omega n$
10.9 ±2.5	BENAKSAS 72C	OSPK	$e^+e^- \rightarrow \pi^0\gamma$
8.1 ±2.0	BALDIN 71	HLBC	$2.9\pi^+\rho$
13 ±4	JACQUET 69B	HLBC	$2.05\pi^+\rho \rightarrow \pi^+\rho\omega$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
9.7 ±0.2 ±0.5	<sup>2,3</sup> ACHASOV 03D	RVUE	$0.44-2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
9.9 ±0.7	<sup>2</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$
<sup>1</sup> From $\omega\pi^0 \rightarrow \pi^0\pi^0\gamma(m_\phi)/\sigma_0^{\omega\pi^0} \rightarrow \pi^+\pi^-\pi^0\pi^0(m_\phi)$ with a phase-space correction factor of 1/1.023.			
<sup>2</sup> Not independent of the corresponding $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .			
<sup>3</sup> Using ACHASOV 03. Based on 1.2M events.			

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}} \qquad \Gamma_3/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.53<math>^{+0.11}_{-0.13}</math> OUR FIT</b>	Error includes scale factor of 1.2.			
<b>1.49±0.13 OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.			
1.46±0.12±0.02	900k	<sup>1</sup> AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
1.30±0.24±0.05	11.2k	<sup>2</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
2.38 $^{+1.77}_{-0.90}$ ±0.18	5.4k	<sup>3</sup> ACHASOV 02E	SND	$1.1-1.38 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
2.3 ±0.5		BARKOV 85	OLYA	$e^+e^- \rightarrow \pi^+\pi^-$
1.6 $^{+0.9}_{-0.7}$		QUENZER 78	DM1	$e^+e^- \rightarrow \pi^+\pi^-$
3.6 ±1.9		BENAKSAS 72	OSPK	$e^+e^- \rightarrow \pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.75±0.11	4.5M	<sup>4</sup> ACHASOV 05A	SND	$e^+e^- \rightarrow \pi^+\pi^-$
2.01±0.29		<sup>5</sup> BENAYOUN 03	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
1.9 ±0.3		<sup>6</sup> GARDNER 99	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
2.3 ±0.4		<sup>7</sup> BENAYOUN 98	RVUE	$e^+e^- \rightarrow \pi^+\pi^-, \mu^+\mu^-$
1.0 ±0.11		<sup>8</sup> WICKLUND 78	ASPK	$3,4,6\pi^\pm N$
1.22±0.30		ALVENSLEB... 71C	CNTR	Photoproduction
1.3 $^{+1.2}_{-0.9}$		MOFFEIT 71	HBC	$2.8,4,7\gamma p$
0.88 $^{+0.28}_{-0.20}$		<sup>9</sup> BIGGS 70B	CNTR	$4.2\gamma C \rightarrow \pi^+\pi^-C$

- <sup>1</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.  
<sup>2</sup> Update of AKHMETSHIN 02.  
<sup>3</sup> From the  $m_{\pi^+\pi^-}$  spectrum taking into account the interference of the  $\rho\pi$  and  $\omega\pi$  amplitudes.  
<sup>4</sup> Using  $\Gamma(\omega \rightarrow e^+e^-)$  from the 2004 Edition of this Review (PDG 04).  
<sup>5</sup> Using the data of AKHMETSHIN 02 in the hidden local symmetry model.  
<sup>6</sup> Using the data of BARKOV 85.  
<sup>7</sup> Using the data of BARKOV 85 in the hidden local symmetry model.  
<sup>8</sup> From a model-dependent analysis assuming complete coherence.  
<sup>9</sup> Re-evaluated under  $\Gamma(\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$  by BEHREND 71 using more accurate  $\omega \rightarrow \rho$  photoproduction cross-section ratio.

WEIGHTED AVERAGE  
1.49±0.13 (Error scaled by 1.3)



### $\Gamma(\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_3/\Gamma_1$

See also  $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0172±0.0014 OUR FIT</b>	Error includes scale factor of 1.2.		
<b>0.026 ±0.005 OUR AVERAGE</b>			
0.021 $\pm 0.028$ $-0.009$	<sup>1,2</sup> RATCLIFF	72 ASPK	15 $\pi^-\rho \rightarrow n2\pi$
0.028 ±0.006	<sup>1</sup> BEHREND	71 ASPK	Photoproduction
0.022 $\pm 0.009$ $-0.01$	<sup>3</sup> ROOS	70 RVUE	

<sup>1</sup> The fitted width of these data is 160 MeV in agreement with present average, thus the  $\omega$  contribution is overestimated. Assuming  $\rho$  width 145 MeV.

<sup>2</sup> Significant interference effect observed. NB of  $\omega \rightarrow 3\pi$  comes from an extrapolation.

<sup>3</sup> ROOS 70 combines ABRAMOVICH 70 and BIZZARRI 70.

### $\Gamma(\pi^+\pi^-)/\Gamma(\pi^0\gamma)$ $\Gamma_3/\Gamma_2$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.20±0.04</b>	1.98M	<sup>1</sup> ALOISIO	03 KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$

<sup>1</sup> Using the data of ALOISIO 02D.

### $\Gamma(\text{neutrals})/\Gamma_{\text{total}}$ $(\Gamma_2+\Gamma_4)/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.091±0.006 OUR FIT</b>				
<b>0.081±0.011 OUR AVERAGE</b>				
0.075±0.025		BIZZARRI	71 HBC	0.0 $p\bar{p}$
0.079±0.019		DEINET	69B OSPK	1.5 $\pi^-\rho$
0.084±0.015		BOLLINI	68C CNTR	2.1 $\pi^-\rho$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.073±0.018	42	BASILE	72B CNTR	1.67 $\pi^-\rho$

### $\Gamma(\text{neutrals})/\Gamma(\pi^+\pi^-\pi^0)$ $(\Gamma_2+\Gamma_4)/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.102±0.008 OUR FIT</b>				
<b>0.103±0.011 OUR AVERAGE</b>				
0.15 ±0.04	46	AGUILAR...	72B HBC	3.9,4.6 $K^-\rho$
0.10 ±0.03	19	BARASH	67B HBC	0.0 $\bar{p}p$
0.134±0.026	850	DIGIUGNO	66B CNTR	1.4 $\pi^-\rho$
0.097±0.016	348	FLATTE	66 HBC	1.4 - 1.7 $K^-\rho \rightarrow \Lambda MM$
0.06 $\pm 0.05$ $-0.02$		JAMES	66 HBC	2.1 $\pi^+\rho$
0.08 ±0.03	35	KRAEMER	64 DBC	1.2 $\pi^+d$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.11 ±0.02	20	BUSCHBECK	63 HBC	1.5 $K^-\rho$

### $\Gamma(\pi^0\gamma)/\Gamma(\text{neutrals})$ $\Gamma_2/(\Gamma_2+\Gamma_4)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.78±0.07		<sup>1</sup> DAKIN	72 OSPK	1.4 $\pi^-\rho \rightarrow nMM$
>0.81	90	DEINET	69B OSPK	

<sup>1</sup> Error statistical only. Authors obtain good fit also assuming  $\pi^0\gamma$  as the only neutral decay.

### $\Gamma(\text{neutrals})/\Gamma(\text{charged particles})$ $(\Gamma_2+\Gamma_4)/(\Gamma_1+\Gamma_3)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.100±0.008 OUR FIT</b>			
<b>0.124±0.021</b>	FELDMAN	67C OSPK	1.2 $\pi^-\rho$

### $\Gamma(\eta\gamma)/\Gamma_{\text{total}}$ $\Gamma_5/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.6 ±0.4 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>6.3 ±1.3 OUR AVERAGE</b>	Error includes scale factor of 1.2.			
6.6 ±1.7		<sup>1</sup> ABELE	97E CBAR	0.0 $\bar{p}p \rightarrow 5\gamma$
8.3 ±2.1		ALDE	93 GAM2	38 $\pi^-\rho \rightarrow \omega n$
3.0 $\pm 2.5$ $-1.8$		<sup>2</sup> ANDREWS	77 CNTR	6.7-10 $\gamma Cu$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.3 ±0.5 ±0.1	33k	<sup>3</sup> ACHASOV	07B SND	0.6-1.38 $e^+e^- \rightarrow \eta\gamma$
4.44 $\pm 2.59$ $-1.83$	17.4k	<sup>4,5</sup> AKHMETSHIN	05 CMD2	0.60-1.38 $e^+e^- \rightarrow \eta\gamma$
5.10±0.72±0.34	23k	<sup>6</sup> AKHMETSHIN	01B CMD2	$e^+e^- \rightarrow \eta\gamma$
0.7 to 5.5		<sup>7</sup> CASE	00 CBAR	0.0 $p\bar{p} \rightarrow \eta\eta\gamma$
6.56 $\pm 2.41$ $-2.55$	3525	<sup>2,8</sup> BENAYOUN	96 RVUE	$e^+e^- \rightarrow \eta\gamma$
7.3 ±2.9		<sup>2,4</sup> DOLINSKY	89 ND	$e^+e^- \rightarrow \eta\gamma$

<sup>1</sup> No flat  $\eta\eta\gamma$  background assumed.

<sup>2</sup> Solution corresponding to constructive  $\omega\rho$  interference.

<sup>3</sup> ACHASOV 07B reports  $[\Gamma(\omega(782) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow e^+e^-)] = (3.10 \pm 0.31 \pm 0.11) \times 10^{-8}$  which we divide by our best value  $B(\omega(782) \rightarrow e^+e^-) = (7.28 \pm 0.14) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>4</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .

<sup>5</sup> Using  $B(\omega \rightarrow e^+e^-) = (7.14 \pm 0.13) \times 10^{-5}$  and  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>6</sup> Using  $B(\omega \rightarrow e^+e^-) = (7.07 \pm 0.19) \times 10^{-5}$  and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ . Solution corresponding to constructive  $\omega\rho$  interference. The combined

fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively). Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .

<sup>7</sup> Depending on the degree of coherence with the flat  $\eta\eta\gamma$  background and using  $B(\omega \rightarrow \pi^0\gamma) = (8.5 \pm 0.5) \times 10^{-2}$ .

<sup>8</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, DOLINSKY 91 taking into account the triangle anomaly contributions.

### $\Gamma(\eta\gamma)/\Gamma(\pi^0\gamma)$ $\Gamma_5/\Gamma_2$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.0098±0.0024	<sup>1</sup> ALDE	93 GAM2	38 $\pi^-\rho \rightarrow \omega n$
0.0082±0.0033	<sup>2</sup> DOLINSKY	89 ND	$e^+e^- \rightarrow \eta\gamma$
0.010 ±0.045	APEL	72B OSPK	4-8 $\pi^-\rho \rightarrow n3\gamma$

<sup>1</sup> Model independent determination.

<sup>2</sup> Solution corresponding to constructive  $\omega\rho$  interference.

### $\Gamma(\pi^0e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_6/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.7 ±0.6 OUR FIT</b>				
<b>7.7 ±0.6 OUR AVERAGE</b>				
7.61±0.53±0.64		ACHASOV	08 SND	0.36-0.97 $e^+e^- \rightarrow \pi^0e^+e^-$
8.19±0.71±0.62		AKHMETSHIN	05A CMD2	0.72-0.84 $e^+e^-$
5.9 ±1.9	43	DOLINSKY	88 ND	$e^+e^- \rightarrow \pi^0e^+e^-$

### $\Gamma(\pi^0\mu^+\mu^-)/\Gamma_{\text{total}}$ $\Gamma_7/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.3 ±0.4 OUR FIT</b>	Error includes scale factor of 2.1.			
<b>1.3 ±0.4 OUR AVERAGE</b>	Error includes scale factor of 2.1.			
1.72±0.25±0.14	3k	ARNALDI	09 NA60	158A In-In collisions
0.96±0.23		DZHELADIN	81B CNTR	25-33 $\pi^-\rho \rightarrow \omega n$

### $\Gamma(\eta e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_8/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<1.1	AKHMETSHIN	05A CMD2	0.72-0.84 $e^+e^-$

### $\Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.728±0.014 OUR FIT</b>	Error includes scale factor of 1.3.			
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.700±0.016	11200	<sup>1,2</sup> AKHMETSHIN	04 CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.752±0.004±0.024	1.2M	<sup>2,3</sup> ACHASOV	03D RVUE	0.44-2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.714±0.036		<sup>2</sup> DOLINSKY	89 ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.72 ±0.03		<sup>2</sup> BARKOV	87 CMD	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.64 ±0.04	1488	<sup>2</sup> KURDADZE	83B OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.675±0.069	433	<sup>2</sup> CORDIER	80 DM1	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.83 ±0.10	451	<sup>2</sup> BENAKSAS	72B OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.77 ±0.06		<sup>4</sup> AUGUSTIN	69D OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.65 ±0.13	33	<sup>5</sup> ASTVACAT...	68 OSPK	Assume SU(3)+mixing

<sup>1</sup> Using  $B(\omega \rightarrow \pi^+\pi^-\pi^0) = 0.891 \pm 0.007$ . Update of AKHMETSHIN 00C.

<sup>2</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}^2$ .

<sup>3</sup> Using ACHASOV 03, ACHASOV 03D and  $B(\omega \rightarrow \pi^+\pi^-) = (1.70 \pm 0.28)\%$ .

<sup>4</sup> Rescaled by us to correspond to  $\omega$  width 8.4 MeV. Systematic errors underestimated.

<sup>5</sup> Not resolved from  $\rho$  decay. Error statistical only.

### $\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2</b>	90	ACHASOV	09A SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<200	90	KURDADZE	86 OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

### $\Gamma(\pi^+\pi^-\gamma)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.0036</b>	95	WEIDENAUER	90 ASTE	$p\bar{p} \rightarrow \pi^+\pi^-\pi^+\pi^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.004	95	BITYUKOV	88B SPEC	32 $\pi^-\rho \rightarrow \pi^+\pi^-\gamma X$

### $\Gamma(\pi^+\pi^-\gamma)/\Gamma(\pi^+\pi^-\pi^0)$ $\Gamma_{11}/\Gamma_1$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.066	90	KALBFLEISCH	75 HBC	2.18 $K^-\rho \rightarrow \Lambda\pi^+\pi^-\gamma$
<0.05	90	FLATTE	66 HBC	1.2-1.7 $K^-\rho \rightarrow \Lambda\pi^+\pi^-\gamma$

### $\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1 × 10<sup>-3</sup></b>	90	KURDADZE	88 OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

# Meson Particle Listings

## $\omega(782)$

$\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>6.6±1.1 OUR FIT</b>					
<b>6.5±1.2 OUR AVERAGE</b>					
$6.4^{+2.4}_{-2.0}\pm 0.8$	190	<sup>1</sup> AKHMETSHIN 04B	CMD2	$0.6\text{--}0.97\ e^+e^- \rightarrow \pi^0\pi^0\gamma$	
$6.6^{+1.4}_{-1.3}\pm 0.6$	295	ACHASOV 02F	SND	$0.36\text{--}0.97\ e^+e^- \rightarrow \pi^0\pi^0\gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$11.8^{+2.1}_{-1.9}\pm 1.4$	190	<sup>2</sup> AKHMETSHIN 04B	CMD2	$0.6\text{--}0.97\ e^+e^- \rightarrow \pi^0\pi^0\gamma$	
$7.8\pm 2.7\pm 2.0$	63	<sup>1,3</sup> ACHASOV 00G	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
$12.7\pm 2.3\pm 2.5$	63	<sup>2,3</sup> ACHASOV 00G	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
<sup>1</sup> In the model assuming the $\rho \rightarrow \pi^0\pi^0\gamma$ decay via the $\omega\pi$ and $f_0(500)\gamma$ mechanisms.					
<sup>2</sup> In the model assuming the $\rho \rightarrow \pi^0\pi^0\gamma$ decay via the $\omega\pi$ mechanism only.					
<sup>3</sup> Superseded by ACHASOV 02F.					

$\Gamma(\pi^0\pi^0\gamma)/\Gamma(\pi^+\pi^-\pi^0)$					$\Gamma_{13}/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.00045</b>	90	DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.08	95	JACQUET 69B	HLBC	$2.05\ \pi^+p \rightarrow \pi^+p\omega$	

$\Gamma(\pi^0\pi^0\gamma)/\Gamma(\pi^0\gamma)$					$\Gamma_{13}/\Gamma_2$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.0±1.3 OUR FIT</b>					
<b>8.5±2.9</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 50	90	DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
<1800	95	KEYNE 76	CNTR	$\pi^-p \rightarrow \omega n$	
<1500	90	BENAKSAS 72c	OSPK	$e^+e^-$	
<1400		BALDIN 71	HLBC	$2.9\ \pi^+p$	
<1000	90	BARMIN 64	HLBC	$1.3\text{--}2.8\ \pi^-p$	

$\Gamma(\pi^0\pi^0\gamma)/\Gamma(\text{neutrals})$					$\Gamma_{13}/(\Gamma_2+\Gamma_4)$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.22\pm 0.07$		<sup>1</sup> DAKIN 72	OSPK	$1.4\ \pi^-p \rightarrow n\text{MM}$	
<0.19	90	DEINET 69B	OSPK		
<sup>1</sup> See $\Gamma(\pi^0\gamma)/\Gamma(\text{neutrals})$ .					

$\Gamma(\eta\pi^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{14}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.3</b>	90	AKHMETSHIN 04B	CMD2	$0.6\text{--}0.97\ e^+e^- \rightarrow \eta\pi^0\gamma$	

$\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>9.0±3.1 OUR FIT</b>					
<b>9.0±2.9±1.1</b>					
	18	HEISTER 02c	ALEP	$Z \rightarrow \mu^+\mu^- + X$	

$\Gamma(\mu^+\mu^-)/\Gamma(\pi^+\pi^-\pi^0)$					$\Gamma_{15}/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.2</b>	90	WILSON 69	OSPK	$12\ \pi^-C \rightarrow \text{Fe}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.7	74	FLATTE 66	HBC	$1.2\text{--}1.7\ K^-p \rightarrow \Lambda\mu^+\mu^-$	
<1.2		BARBARO-... 65	HBC	$2.7\ K^-p$	

$\Gamma(\pi^0\mu^+\mu^-)/\Gamma(\mu^+\mu^-)$					$\Gamma_7/\Gamma_{15}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.2\pm 0.6$	30	<sup>1</sup> DZHELYADIN 79	CNTR	$25\text{--}33\ \pi^-p$	
<sup>1</sup> Superseded by DZHELYADIN 81B result above.					

$\Gamma(3\gamma)/\Gamma_{\text{total}}$					$\Gamma_{16}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.9</b>	95	<sup>1</sup> ABELE 97E	CBAR	$0.0\ \overline{p}p \rightarrow 5\gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<2	90	<sup>1</sup> PROKOSHKIN 95	GAM2	$38\ \pi^-p \rightarrow 3\gamma n$	
<sup>1</sup> From direct $3\gamma$ decay search.					

$\Gamma(\eta\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
Violates C conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.001	90	ALDE 94B	GAM2	$38\pi^-p \rightarrow \eta\pi^0 n$	

$[\Gamma(\eta\gamma) + \Gamma(\eta\pi^0)]/\Gamma(\pi^+\pi^-\pi^0)$					$(\Gamma_5+\Gamma_{17})/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.016</b>	90	<sup>1</sup> FLATTE 66	HBC	$1.2\text{--}1.7\ K^-p \rightarrow \Lambda\pi^+\pi^-\text{MM}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.045	95	JACQUET 69B	HLBC	$2.05\ \pi^+p \rightarrow \pi^+p\omega$	
<sup>1</sup> Restated by us using $B(\eta \rightarrow \text{charged modes}) = 29.2\%$ .					

$\Gamma(\eta\pi^0)/\Gamma(\pi^0\gamma)$					$\Gamma_{17}/\Gamma_2$
Violates C conservation.					
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.6</b>	90	<sup>1</sup> STAROSTIN 09	CRYM	$\gamma p \rightarrow \eta\pi^0 p$	
<sup>1</sup> STAROSTIN 09 reports $[\Gamma(\omega(782) \rightarrow \eta\pi^0)/\Gamma(\omega(782) \rightarrow \pi^0\gamma)] \times [B(\eta \rightarrow 2\gamma)] < 1.01 \times 10^{-3}$ which we divide by our best value $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .					

$\Gamma(2\pi^0)/\Gamma(\pi^0\gamma)$					$\Gamma_{18}/\Gamma_2$
Violates C conservation and Bose-Einstein statistics.					
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.59</b>	90	STAROSTIN 09	CRYM	$\gamma p \rightarrow 2\pi^0 p$	

$\Gamma(3\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
Violates C conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< $3 \times 10^{-4}$	90	PROKOSHKIN 95	GAM2	$38\ \pi^-p \rightarrow 3\pi^0 n$	

$\Gamma(3\pi^0)/\Gamma(\pi^0\gamma)$					$\Gamma_{19}/\Gamma_2$
Violates C conservation.					
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.72</b>	90	STAROSTIN 09	CRYM	$\gamma p \rightarrow 3\pi^0 p$	

$\Gamma(3\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$					$\Gamma_{19}/\Gamma_1$
Violates C conservation.					
VALUE	CL%	DOCUMENT ID	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.009	90	BARBERIS 01	$450\ p\overline{p} \rightarrow p_f 3\pi^0 p_s$		

### PARAMETER $\Lambda$ IN $\omega \rightarrow \pi^0\mu^+\mu^-$ DECAY

In the pole approximation the electromagnetic transition form factor for a resonance of mass  $M$  is given by the expression:  
 $|F|^2 = (1 - M^2/\Lambda^2)^{-2}$ ,  
where for the parameter  $\Lambda$  vector dominance predicts  $\Lambda = M_\rho \approx 0.770$  GeV. The ARNALDI 09 measurement is in obvious conflict with this expectation. Note that for  $\eta \rightarrow \mu^+\mu^-\gamma$  decay ARNALDI 09 and DZHELYADIN 80 obtain the value of  $\Lambda$  consistent with vector dominance.

VALUE (GeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.668±0.009±0.003</b>	3k	ARNALDI 09	NA60	158A In-In collisions
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.65 \pm 0.03$		DZHELYADIN 81B	CNTR	$25\text{--}33\ \pi^-p \rightarrow \omega n$

### $\omega(782)$ REFERENCES

ACHASOV 13	PR D88 054013	M.N. Achasov <i>et al.</i>	(SND Collab.)
BENAYOUN 13	EPJ C73 2453	M. Benayoun, P. David, L. DelBuono (PARIN, BERLIN+)	
DAVIER 13	EPJ C73 2597	M. Davier <i>et al.</i>	
LEES 12G	PR D86 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
NIECKNIG 12	EPJ C72 2014	F. Niecknig, B. Kubis, S.P. Schneider	(BONN)
BENAYOUN 10	EPJ C65 211	M. Benayoun <i>et al.</i>	
ACHASOV 09A	JETP 109 379	M.N. Achasov <i>et al.</i>	(SND Collab.)
Translated from ZETF 136 442.			
ARNALDI 09	PL B677 260	R. Arnaldi <i>et al.</i>	(NA60 Collab.)
STAROSTIN 09	PL C79 065201	A. Starostin <i>et al.</i>	(Crystal Ball Collab. at MAMI)
ACHASOV 08	JETP 107 61	M.N. Achasov <i>et al.</i>	(SND Collab.)
Translated from ZETF 134 80.			
AMBROSINO 08G	PL B669 223	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
ACHASOV 07B	PL D76 077101	M.N. Achasov <i>et al.</i>	(SND Collab.)
AKHMETSHIN 07	PL B648 28	R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ACHASOV 06	JETP 103 380	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
Translated from ZETF 130 437.			





# Meson Particle Listings

## $\eta'(958)$

Charge conjugation (C), Parity (P), Lepton family number (LF) violating modes				
$\Gamma_{25}$	$\pi^+ \pi^-$	$P, CP$	$< 6$	$\times 10^{-5}$ 90%
$\Gamma_{26}$	$\pi^0 \pi^0$	$P, CP$	$< 4$	$\times 10^{-4}$ 90%
$\Gamma_{27}$	$\pi^0 e^+ e^-$	$C$	$[a] < 1.4$	$\times 10^{-3}$ 90%
$\Gamma_{28}$	$\eta e^+ e^-$	$C$	$[a] < 2.4$	$\times 10^{-3}$ 90%
$\Gamma_{29}$	$3\gamma$	$C$	$< 1.0$	$\times 10^{-4}$ 90%
$\Gamma_{30}$	$\mu^+ \mu^- \pi^0$	$C$	$[a] < 6.0$	$\times 10^{-5}$ 90%
$\Gamma_{31}$	$\mu^+ \mu^- \eta$	$C$	$[a] < 1.5$	$\times 10^{-5}$ 90%
$\Gamma_{32}$	$e \mu$	$LF$	$< 4.7$	$\times 10^{-4}$ 90%

[a] C parity forbids this to occur as a single-photon process.

### CONSTRAINED FIT INFORMATION

An overall fit to the total width, a partial width, 2 combinations of partial widths obtained from integrated cross section, and 16 branching ratios uses 46 measurements and one constraint to determine 9 parameters. The overall fit has a  $\chi^2 = 52.8$  for 38 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−2							
$x_3$	−77	−58						
$x_4$	−11	−13	2					
$x_6$	−29	−25	32	−1				
$x_7$	−24	−19	30	0	9			
$x_{10}$	0	−2	−2	0	−1	−1		
$x_{18}$	−4	−6	−5	−1	−3	−2	0	
$\Gamma$	25	5	−19	3	−71	−5	1	3
	$x_1$	$x_2$	$x_3$	$x_4$	$x_6$	$x_7$	$x_{10}$	$x_{18}$

Mode	Rate (MeV)
$\Gamma_1$ $\pi^+ \pi^- \eta$	0.085 ± 0.004
$\Gamma_2$ $\rho^0 \gamma$ (including non-resonant $\pi^+ \pi^- \gamma$ )	0.0574 ± 0.0028
$\Gamma_3$ $\pi^0 \pi^0 \eta$	0.0440 ± 0.0023
$\Gamma_4$ $\omega \gamma$	0.00517 ± 0.00035
$\Gamma_6$ $\gamma \gamma$	0.00435 ± 0.00013
$\Gamma_7$ $3\pi^0$	(4.3 ± 0.4) $\times 10^{-4}$
$\Gamma_{10}$ $\pi^+ \pi^- \pi^0$	(7.5 ± 0.8) $\times 10^{-4}$
$\Gamma_{18}$ $\pi^+ \pi^- e^+ e^-$	(4.7 $\begin{smallmatrix} +2.6 \\ -1.9 \end{smallmatrix}$ ) $\times 10^{-4}$

### $\eta'(958)$ PARTIAL WIDTHS

$\Gamma(\gamma\gamma)$				$\Gamma_6$
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.35 ± 0.14 OUR FIT</b>				
<b>4.28 ± 0.19 OUR AVERAGE</b>				
4.17 ± 0.10 ± 0.27	2000	<sup>1</sup> ACCIARRI	98Q L3	$e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \gamma$
4.53 ± 0.29 ± 0.51	266	KARCH	92 CBAL	$e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$
3.61 ± 0.13 ± 0.48		<sup>2</sup> BEHREND	91 CELL	$e^+ e^- \rightarrow e^+ e^- \eta'(958)$
4.6 ± 1.1 ± 0.6	23	BARU	90 MD1	$e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \gamma$
4.57 ± 0.25 ± 0.44		BUTLER	90 MRK2	$e^+ e^- \rightarrow e^+ e^- \eta'(958)$
5.08 ± 0.24 ± 0.71	547	<sup>3</sup> ROE	90 ASP	$e^+ e^- \rightarrow e^+ e^- 2\gamma$
3.8 ± 0.7 ± 0.6	34	AIHARA	88c TPC	$e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$
4.9 ± 0.5 ± 0.5	136	<sup>4</sup> WILLIAMS	88 CBAL	$e^+ e^- \rightarrow e^+ e^- 2\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.7 ± 0.6 ± 0.9	143	<sup>5</sup> GIDAL	87 MRK2	$e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$
4.0 ± 0.9		<sup>6</sup> BARTEL	85E JADE	$e^+ e^- \rightarrow e^+ e^- 2\gamma$

- <sup>1</sup> No non-resonant  $\pi^+ \pi^-$  contribution found.  
<sup>2</sup> Reevaluated by us using  $B(\eta' \rightarrow \rho(770) \gamma) = (30.2 \pm 1.3)\%$ .  
<sup>3</sup> Reevaluated by us using  $B(\eta' \rightarrow \gamma \gamma) = (2.11 \pm 0.13)\%$ .  
<sup>4</sup> Reevaluated by us using  $B(\eta' \rightarrow \gamma \gamma) = (2.11 \pm 0.13)\%$ .  
<sup>5</sup> Superseded by BUTLER 90.  
<sup>6</sup> Systematic error not evaluated.

$\Gamma(e^+ e^-)$				$\Gamma_{23}$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.1 <math>\times 10^{-3}</math></b>	90	<sup>1,2</sup> ACHASOV	15 SND	$0.958 e^+ e^- \rightarrow \pi \pi \eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 2.0 $\times 10^{-3}$	90	<sup>2</sup> ACHASOV	15 SND	$0.958 e^+ e^- \rightarrow \pi \pi \eta$
< 2.4 $\times 10^{-3}$	90	<sup>2</sup> AKHMETSHIN	15 CMD3	$0.958 e^+ e^- \rightarrow \pi^+ \pi^- \eta$

- <sup>1</sup> Combining data of ACHASOV 15 and AKHMETSHIN 15.  
<sup>2</sup> Using  $\eta$  and  $\eta'$  branching fractions from PDG 14.

### $\eta'(958) \Gamma(i) \Gamma(\gamma\gamma) / \Gamma(\text{total})$

This combination of a partial width with the partial width into  $\gamma\gamma$  and with the total width is obtained from the integrated cross section into channel(i) in the  $\gamma\gamma$  annihilation.

$\Gamma(\gamma\gamma) \times \Gamma(\rho^0 \gamma \text{ (including non-resonant } \pi^+ \pi^- \gamma)) / \Gamma_{\text{total}}$				$\Gamma_6 \Gamma_2 / \Gamma$
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.27 ± 0.04 OUR FIT</b>				
<b>1.26 ± 0.07 OUR AVERAGE</b>				Error includes scale factor of 1.2.
1.09 ± 0.04 ± 0.13		BEHREND	91 CELL	$e^+ e^- \rightarrow e^+ e^- \rho(770)^0 \gamma$
1.35 ± 0.09 ± 0.21		AIHARA	87 TPC	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$
1.13 ± 0.04 ± 0.13	867	ALBRECHT	87B ARG	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$
1.53 ± 0.09 ± 0.21		ALTHOFF	84E TASS	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$
1.14 ± 0.08 ± 0.11	243	BERGER	84B PLUT	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$
1.73 ± 0.34 ± 0.35	95	JENNI	83 MRK2	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$
1.49 ± 0.13 ± 0.027	213	BARTEL	82B JADE	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.85 ± 0.31 ± 0.24	43	BEHREND	83B CELL	$e^+ e^- \rightarrow e^+ e^- \rho \gamma$

$\Gamma(\gamma\gamma) \times \Gamma(\pi^0 \pi^0 \eta) / \Gamma_{\text{total}}$				$\Gamma_6 \Gamma_3 / \Gamma$
VALUE (keV)		DOCUMENT ID	TECN	COMMENT
<b>0.97 ± 0.05 OUR FIT</b>				
<b>0.92 ± 0.06 ± 0.11</b>		<sup>1</sup> KARCH	92 CBAL	$e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.95 ± 0.05 ± 0.08		<sup>2</sup> KARCH	90 CBAL	$e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$
1.00 ± 0.08 ± 0.10		<sup>2,3</sup> ANTREASNYAN	87 CBAL	$e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$
<sup>1</sup> Reevaluated by us using $B(\eta \rightarrow \gamma \gamma) = (39.21 \pm 0.34)\%$ . Supersedes ANTREASNYAN 87 and KARCH 90. <sup>2</sup> Superseded by KARCH 92. <sup>3</sup> Using $BR(\eta \rightarrow 2\gamma) = (38.9 \pm 0.5)\%$ .				

### $\eta'(958) \Gamma(i) \Gamma(e^+ e^-) / \Gamma(\text{total})$

$\Gamma(\pi^+ \pi^- \eta) \times \Gamma(e^+ e^-) / \Gamma_{\text{total}}$				$\Gamma_1 \Gamma_{23} / \Gamma$
VALUE ( $10^{-3}$ eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.0</b>	90	<sup>1</sup> AKHMETSHIN	15 CMD3	$0.958 e^+ e^- \rightarrow \pi^+ \pi^- \eta$
<sup>1</sup> AKHMETSHIN 15 reports $[\Gamma(\eta'(958) \rightarrow \pi^+ \pi^- \eta) \times \Gamma(\eta'(958) \rightarrow e^+ e^-) / \Gamma_{\text{total}}] \times [B(\eta \rightarrow 2\gamma)] < 4.1 \times 10^{-4}$ eV which we divide by our best value $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .				

### $\eta'(958)$ BRANCHING RATIOS

$\Gamma(\pi^+ \pi^- \eta) / \Gamma_{\text{total}}$				$\Gamma_1 / \Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.429 ± 0.007 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.424 ± 0.011 ± 0.004	1.2k	<sup>1</sup> PEDLAR	09 CLEO	$J/\psi \rightarrow \gamma \eta'$
<sup>1</sup> Not independent of other $\eta'$ branching fractions and ratios in PEDLAR 09.				

$\Gamma(\pi^+ \pi^- \eta \text{ (charged decay)}) / \Gamma_{\text{total}}$				0.286 $\Gamma_1 / \Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1228 ± 0.0020 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.123 ± 0.014	107	RITTENBERG	69 HBC	$1.7\text{--}2.7 K^- p$
0.10 ± 0.04	10	LONDON	66 HBC	$2.24 K^- p \rightarrow \Lambda 2\pi^+ 2\pi^- \pi^0$
0.07 ± 0.04	7	BADIER	65B HBC	$3 K^- p$

$\Gamma(\pi^+ \pi^- \eta \text{ (neutral decay)}) / \Gamma_{\text{total}}$				0.714 $\Gamma_1 / \Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.307 ± 0.005 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.314 ± 0.026	281	RITTENBERG	69 HBC	$1.7\text{--}2.7 K^- p$

$\Gamma(\rho^0 \gamma \text{ (including non-resonant } \pi^+ \pi^- \gamma)) / \Gamma_{\text{total}}$				$\Gamma_2 / \Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.291 ± 0.005 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.287 ± 0.007 ± 0.004	0.2k	<sup>1</sup> PEDLAR	09 CLEO	$J/\psi \rightarrow \gamma \eta'$
0.329 ± 0.033	298	RITTENBERG	69 HBC	$1.7\text{--}2.7 K^- p$
0.2 ± 0.1	20	LONDON	66 HBC	$2.24 K^- p \rightarrow \Lambda \pi^+ \pi^- \gamma$
0.34 ± 0.09	35	BADIER	65B HBC	$3 K^- p$
<sup>1</sup> Not independent of other $\eta'$ branching fractions and ratios in PEDLAR 09.				

$\Gamma(\rho^0 \gamma \text{ (including non-resonant } \pi^+ \pi^- \gamma)) / \Gamma(\pi^+ \pi^- \eta)$				$\Gamma_2 / \Gamma_1$
VALUE		DOCUMENT ID	TECN	COMMENT
<b>0.678 ± 0.017 OUR FIT</b>				
<b>0.683 ± 0.020 OUR AVERAGE</b>				
0.677 ± 0.024 ± 0.011		PEDLAR	09 CLE3	$J/\psi \rightarrow \eta' \gamma$
0.69 ± 0.03		ABLIKIM	06E BES2	$J/\psi \rightarrow \eta' \gamma$

See key on page 601

## Meson Particle Listings

 $\eta'(958)$ 
 $\Gamma(\rho^0\gamma(\text{including non-resonant } \pi^+\pi^-\gamma))/\Gamma(\pi^+\pi^-\eta(\text{neutral decay}))/\Gamma_2/0.714\Gamma_1$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.950±0.024 OUR FIT</b>				
<b>0.97 ±0.09 OUR AVERAGE</b>				
0.70 ±0.22		AMSLER	04B	CBAR $0\bar{p}p \rightarrow \pi^+\pi^-\eta$
1.07 ±0.17		BELADIDZE	92c	VES $36\pi^-\text{Be} \rightarrow \pi^-\eta'\eta\text{Be}$
0.92 ±0.14	473	DANBURG	73	HBC $2.2K^-p \rightarrow \Lambda X^0$
1.11 ±0.18	192	JACOBS	73	HBC $2.9K^-p \rightarrow \Lambda X^0$

 $\Gamma(\pi^0\pi^0\eta)/\Gamma_{\text{total}} \quad \Gamma_3/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.223±0.008 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.235±0.013±0.004	3.2k	<sup>1</sup> PEDLAR	09	CLEO $J/\psi \rightarrow \gamma\eta'$
<sup>1</sup> Not independent of other $\eta'$ branching fractions and ratios in PEDLAR 09.				

 $\Gamma(\pi^0\pi^0\eta(3\pi^0\text{decay}))/\Gamma_{\text{total}} \quad 0.321\Gamma_3/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0716±0.0026 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.11 ±0.06	4	BENSINGER	70	DBC $2.2\pi^+d$

 $\Gamma(\pi^0\pi^0\eta)/\Gamma(\pi^+\pi^-\eta) \quad \Gamma_3/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.519±0.026 OUR FIT</b>			
<b>0.555±0.043±0.013</b>	PEDLAR	09	CLE3 $J/\psi \rightarrow \eta'\gamma$

 $\Gamma(\rho^0\gamma(\text{including non-resonant } \pi^+\pi^-\gamma))/\Gamma(\pi\pi\eta) \quad \Gamma_2/(\Gamma_1+\Gamma_3)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.446±0.012 OUR FIT</b>			
<b>0.43 ±0.02 ±0.02</b>	BARBERIS	98c	OMEG $450pp \rightarrow p_f\eta'p_s$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.31 ±0.15	DAVIS	68	HBC $5.5K^-p$

 $\Gamma(\omega\gamma)/\Gamma_{\text{total}} \quad \Gamma_4/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.62±0.13 OUR FIT</b>				
<b>2.55±0.03±0.16</b>	33.2k	<sup>1</sup> ABLIKIM	15AD	BES3 $J/\psi \rightarrow \eta'\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.34±0.30±0.04	70	<sup>2</sup> PEDLAR	09	CLEO $J/\psi \rightarrow \eta'\eta'$
<sup>1</sup> Using $B(J/\psi \rightarrow \eta'\gamma) = (5.15 \pm 0.16) \times 10^{-3}$ and $B(\omega \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7)\%$ .				
<sup>2</sup> Not independent of other $\eta'$ branching fractions and ratios in PEDLAR 09.				

 $\Gamma(\omega\gamma)/\Gamma(\pi^+\pi^-\eta) \quad \Gamma_4/\Gamma_1$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0610±0.0033 OUR FIT</b>				
<b>0.055 ±0.007 ±0.001</b>		PEDLAR	09	CLE3 $J/\psi \rightarrow \eta'\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.068 ±0.013	68	ZANFINO	77	ASPK $8.4\pi^-p$

 $\Gamma(\omega\gamma)/\Gamma(\pi^0\pi^0\eta) \quad \Gamma_4/\Gamma_3$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.117±0.007 OUR FIT</b>			
<b>0.147±0.016</b>	ALDE	87B	GAM2 $38\pi^-p \rightarrow n4\gamma$

 $\Gamma(\omega e^+e^-)/\Gamma(\omega\gamma) \quad \Gamma_5/\Gamma_4$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
7.71±1.34±0.54	<sup>1</sup> ABLIKIM	15AD	BES3 $J/\psi \rightarrow \eta'\gamma$
<sup>1</sup> Obtained from other ABLIKIM 15AD measurements with common systematics taken into account.			

 $\Gamma(\omega e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_5/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.97±0.34±0.17</b>	66	<sup>1</sup> ABLIKIM	15AD	BES3 $J/\psi \rightarrow \eta'\gamma$
<sup>1</sup> Using $B(J/\psi \rightarrow \eta'\gamma) = (5.15 \pm 0.16) \times 10^{-3}$ and $B(\omega \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7)\%$ .				

 $\Gamma(\rho^0\gamma(\text{including non-resonant } \pi^+\pi^-\gamma))/[\Gamma(\pi^+\pi^-\eta) + \Gamma(\pi^0\pi^0\eta) + \Gamma(\omega\gamma)] \quad \Gamma_2/(\Gamma_1+\Gamma_3+\Gamma_4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.429±0.011 OUR FIT</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.25 ±0.14	DAUBER	64	HBC $1.95K^-p$

 $[\Gamma(\pi^0\pi^0\eta(\text{charged decay})) + \Gamma(\omega(\text{charged decay})\gamma)]/\Gamma_{\text{total}} \quad (0.286\Gamma_3+0.89\Gamma_4)/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0871±0.0026 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.045 ±0.029	42	RITTENBERG	69	HBC $1.7\text{--}2.7K^-p$

 $\Gamma(\pi^+\pi^-\text{neutrals})/\Gamma_{\text{total}} \quad (0.714\Gamma_1+0.286\Gamma_3+0.89\Gamma_4)/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.394±0.004 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.4 ±0.1	39	LONDON	66	HBC $2.24K^-p \rightarrow \Lambda\pi^+\pi^-\text{neutrals}$
0.35 ±0.06	33	BADIER	65B	HBC $3K^-p$

 $\Gamma(\gamma\gamma)/\Gamma_{\text{total}} \quad \Gamma_6/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.21±0.08 OUR FIT</b>				
<b>2.00±0.15 OUR AVERAGE</b>				
1.98 <sup>+0.31</sup> <sub>-0.27</sub> ±0.07	114	<sup>1</sup> WICHT	08	BELL $B^\pm \rightarrow K^\pm\gamma\gamma$
2.00±0.18		<sup>2</sup> STANTON	80	SPEC $8.45\pi^-p \rightarrow n\pi^+\pi^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.25±0.16±0.03	0.3k	<sup>3</sup> PEDLAR	09	CLEO $J/\psi \rightarrow \gamma\eta'$
1.8 ±0.2	6000	<sup>4</sup> APEL	79	NICE $15\text{--}40\pi^-p \rightarrow n2\gamma$
2.5 ±0.7		DUANE	74	MMS $\pi^-p \rightarrow n\text{MM}$
1.71±0.33	68	DALPIAZ	72	CNTR $1.6\pi^-p \rightarrow nX^0$
2.0 <sup>+0.8</sup> <sub>-0.6</sub>	31	HARVEY	71	OSPK $3.65\pi^-p \rightarrow nX^0$

<sup>1</sup> WICHT 08 reports  $[\Gamma(\eta'(958) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow \eta'K^+)] = (1.40<sup>+0.16+0.15</sup><sub>-0.15-0.12</sub>) \times 10^{-6}$  which we divide by our best value  $B(B^+ \rightarrow \eta'K^+) = (7.06 \pm 0.25) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Includes APEL 79 result.

<sup>3</sup> Not independent of other  $\eta'$  branching fractions and ratios in PEDLAR 09.

<sup>4</sup> Data is included in STANTON 80 evaluation.

 $\Gamma(\gamma\gamma)/\Gamma(\pi^+\pi^-\eta) \quad \Gamma_6/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0514±0.0022 OUR FIT</b>			
<b>0.053 ±0.004 ±0.001</b>	PEDLAR	09	CLE3 $J/\psi \rightarrow \eta'\gamma$

 $\Gamma(\gamma\gamma)/\Gamma(\rho^0\gamma(\text{including non-resonant } \pi^+\pi^-\gamma)) \quad \Gamma_6/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0758±0.0033 OUR FIT</b>			
<b>0.080 ±0.008</b>	ABLIKIM	06E	BES2 $J/\psi \rightarrow \eta'\gamma$

 $\Gamma(\gamma\gamma)/\Gamma(\pi^0\pi^0\eta) \quad \Gamma_6/\Gamma_3$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.099±0.004 OUR FIT</b>			
<b>0.105±0.010 OUR AVERAGE</b>	Error includes scale factor of 1.9.		
0.091±0.009	AMSLER	93	CBAR $0.0\bar{p}p$
0.112±0.002±0.006	ALDE	87B	GAM2 $38\pi^-p \rightarrow n2\gamma$

 $\Gamma(\gamma\gamma)/\Gamma(\pi^0\pi^0\eta(\text{neutral decay})) \quad \Gamma_6/0.714\Gamma_3$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.139±0.006 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.188±0.058	16	APEL	72	OSPK $3.8\pi^-p \rightarrow nX^0$

 $\Gamma(\text{neutrals})/\Gamma_{\text{total}} \quad (0.714\Gamma_3+0.09\Gamma_4+\Gamma_6)/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.184±0.006 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.185±0.022	535	BASILE	71	CNTR $1.6\pi^-p \rightarrow nX^0$
0.189±0.026	123	RITTENBERG	69	HBC $1.7\text{--}2.7K^-p$

 $\Gamma(3\pi^0)/\Gamma_{\text{total}} \quad \Gamma_7/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.20±0.20 OUR FIT</b>				
<b>3.7 ±0.4 OUR AVERAGE</b>				
4.79±0.59±1.14	183	<sup>1</sup> ABLIKIM	15P	BES3 $J/\psi \rightarrow K^+K^-3\pi$
3.56±0.22±0.34	309	ABLIKIM	12E	BES3 $J/\psi \rightarrow \gamma(3\pi^0)$

<sup>1</sup> We have added all systematic uncertainties in quadrature to a single value.

 $\Gamma(3\pi^0)/\Gamma(\pi^0\pi^0\eta) \quad \Gamma_7/\Gamma_3$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>99±9 OUR FIT</b>				
<b>78±10 OUR AVERAGE</b>				
86±19	235	BLIK	08	GAMS $32\pi^-p \rightarrow \eta'n$
74±15		ALDE	87B	GAM2 $38\pi^-p \rightarrow n6\gamma$
75±18		BINON	84	GAM2 $30\text{--}40\pi^-p \rightarrow n6\gamma$

 $\Gamma(\mu^+\mu^-\gamma)/\Gamma(\gamma\gamma) \quad \Gamma_8/\Gamma_6$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.9±1.2</b>	33	VIKTOROV	80	CNTR $25,33\pi^-p \rightarrow 2\mu\gamma$

# Meson Particle Listings

## $\eta'(958)$

$\Gamma(\pi^+\pi^-\mu^+\mu^-)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
VALUE (units $10^{-4}$ )	CL%

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.29	90	<sup>1</sup> ABLIKIM	13o	BES3	$J/\psi \rightarrow \gamma\eta'$
<2.4	90	<sup>2</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$

<sup>1</sup> Using  $\Gamma_2/\Gamma = (29.3 \pm 0.6)\%$  from PDG 12.

<sup>2</sup> Not independent of measured value of  $\Gamma_9/\Gamma_1$  from NAIK 09.

$\Gamma(\pi^+\pi^-\mu^+\mu^-)/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_9/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%

<0.5	90	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\mu^+\mu^-)/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] < 1.3 \times 10^{-3}$  which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .

$\Gamma(\pi^+\pi^-\mu^+\mu^-)/\Gamma(\rho^0\gamma(\text{including non-resonant } \pi^+\pi^-\gamma))$	$\Gamma_9/\Gamma_2$
VALUE (units $10^{-4}$ )	CL%

<1.0	90	ABLIKIM	13o	BES3	$J/\psi \rightarrow \gamma\eta'$
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$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS

**3.82±0.35 OUR FIT**

**3.9 ±0.4 OUR AVERAGE**

4.28±0.49±1.11	78	<sup>1</sup> ABLIKIM	15P	BES3	$J/\psi \rightarrow K^+K^-\pi$
3.83±0.15±0.39	1014	ABLIKIM	12E	BES3	$J/\psi \rightarrow \gamma(\pi^+\pi^-\pi^0)$
3.7 <sub>-0.9</sub> <sup>+1.1</sup> ±0.4		<sup>2</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$

<sup>1</sup> We have added all systematic uncertainties in quadrature to a single value.

<sup>2</sup> Not independent of measured value of  $\Gamma_{10}/\Gamma_1$  from NAIK 09.

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{10}/\Gamma_1$
VALUE (units $10^{-3}$ )	EVTS

**8.9 ±0.8 OUR FIT**

8.28 <sub>-2.12</sub> <sup>+2.49</sup> ±0.04	20	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\pi^0)/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] = (21<sub>-5</sub><sup>+6</sup>±2)×10<sup>-3</sup> which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.$

$\Gamma(\pi^0\rho^0)/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$
VALUE	CL%

<0.04	90	RITTENBERG	65	HBC	2.7 $K^-p$
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$\Gamma(2(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$
VALUE (units $10^{-5}$ )	CL%

8.5±0.9±0.3	199	<sup>1</sup> ABLIKIM	14M	BES3	$J/\psi \rightarrow \gamma\eta'$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

< 24	90	<sup>2</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
<100	90	RITTENBERG	69	HBC	1.7-2.7 $K^-p$

<sup>1</sup> ABLIKIM 14M reports  $[\Gamma(\eta'(958) \rightarrow 2(\pi^+\pi^-))/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta'(958))]$  = (4.40 ± 0.35 ± 0.30) × 10<sup>-7</sup> which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta'(958)) = (5.15 \pm 0.16) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Not independent of measured value of  $\Gamma_{12}/\Gamma_1$  from NAIK 09.

$\Gamma(2(\pi^+\pi^-))/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{12}/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%

<0.6	90	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow 2(\pi^+\pi^-))/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] < 1.4 \times 10^{-3}$  which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$
VALUE (units $10^{-4}$ )	CL%

1.8±0.4±0.1	84	<sup>1</sup> ABLIKIM	14M	BES3	$J/\psi \rightarrow \gamma\eta'$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<27	90	<sup>2</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> ABLIKIM 14M reports  $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\pi^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta'(958))]$  = (9.38±1.79±0.89)×10<sup>-7</sup> which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta'(958)) = (5.15 \pm 0.16) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Not independent of measured value of  $\Gamma_{13}/\Gamma_1$  from NAIK 09.

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{13}/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%

<6	90	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\pi^0)/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] < 15 \times 10^{-3}$  which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .

$\Gamma(2(\pi^+\pi^-) \text{ neutrals})/\Gamma_{\text{total}}$	$\Gamma_{14}/\Gamma$
VALUE	CL%

<0.01	95	DANBURG	73	HBC	2.2 $K^-p \rightarrow \Lambda X^0$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.01	90	RITTENBERG	69	HBC	1.7-2.7 $K^-p$
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$\Gamma(2(\pi^+\pi^-)\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{15}/\Gamma$
VALUE	CL%

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.002	90	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
<0.01	90	RITTENBERG	69	HBC	1.7-2.7 $K^-p$

<sup>1</sup> Not independent of measured value of  $\Gamma_{15}/\Gamma_1$  from NAIK 09.

$\Gamma(2(\pi^+\pi^-)\pi^0)/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{15}/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%

<4	90	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow 2(\pi^+\pi^-)\pi^0)/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] < 11 \times 10^{-3}$  which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .

$\Gamma(2(\pi^+\pi^-)2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{16}/\Gamma$
VALUE	CL%

<0.01	95	KALBFLEISCH	64B	HBC	$K^-p \rightarrow \Lambda 2(\pi^+\pi^-)+MM$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.01	90	LONDON	66	HBC	Compilation
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$\Gamma(3(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_{17}/\Gamma$
VALUE (units $10^{-5}$ )	CL%

< 3.1	90	<sup>1</sup> ABLIKIM	13U	BES3	$J/\psi \rightarrow \gamma 3(\pi^+\pi^-)$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

< 53	90	<sup>2</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
<500	95	KALBFLEISCH	64B	HBC	$K^-p \rightarrow \Lambda 2(\pi^+\pi^-)$

<sup>1</sup> Using  $B(J/\psi \rightarrow \gamma\eta'(958)) = (5.16 \pm 0.15) \times 10^{-3}$ .

<sup>2</sup> Not independent of measured value of  $\Gamma_{17}/\Gamma_1$  from NAIK 09.

$\Gamma(3(\pi^+\pi^-))/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{17}/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%

<1.2	90	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow 3(\pi^+\pi^-))/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] < 3.0 \times 10^{-3}$  which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .

$\Gamma(\pi^+\pi^-\text{e}^+\text{e}^-)/\Gamma_{\text{total}}$	$\Gamma_{18}/\Gamma$
VALUE (units $10^{-3}$ )	CL%

2.4 <sub>-1.0</sub> <sup>+1.3</sup>					
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**2.4 ±1.3 OUR FIT**

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.11±0.12±0.14	429	<sup>1</sup> ABLIKIM	13o	BES3	$J/\psi \rightarrow \gamma\eta'$
2.5 <sub>-0.9</sub> <sup>+1.2</sup> ±0.5		<sup>2</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
<6	90	RITTENBERG	65	HBC	2.7 $K^-p$

<sup>1</sup> Using  $\Gamma_2/\Gamma = (29.3 \pm 0.6)\%$  from PDG 12.

<sup>2</sup> Not independent of measured value of  $\Gamma_{18}/\Gamma_1$  from NAIK 09.

$\Gamma(\pi^+\pi^-\text{e}^+\text{e}^-)/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{18}/\Gamma_1$
VALUE (units $10^{-3}$ )	EVTS

5.6 <sub>-2.2</sub> <sup>+3.0</sup>					
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**5.6 ±3.0 OUR FIT**

5.52 <sub>-2.30</sub> <sup>+3.00</sup> ±0.03	8	<sup>1</sup> NAIK	09	CLEO	$J/\psi \rightarrow \gamma\eta'$
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<sup>1</sup> NAIK 09 reports  $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\text{e}^+\text{e}^-)/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] = (14<sub>-5</sub><sup>+7</sup>±3) × 10<sup>-3</sup> which we multiply by our best value  $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.$

$\Gamma(\pi^+\pi^-\text{e}^+\text{e}^-)/\Gamma(\rho^0\gamma(\text{including non-resonant } \pi^+\pi^-\gamma))$	$\Gamma_{18}/\Gamma_2$
VALUE (units $10^{-2}$ )	EVTS

7.2±0.4±0.5	429	ABLIKIM	13o	BES3	$J/\psi \rightarrow \gamma\eta'$
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$\Gamma(\pi^+\text{e}^-\nu_e + \text{c.c.})/\Gamma(\pi^+\pi^-\eta)$	$\Gamma_{19}/\Gamma_1$
VALUE (units $10^{-4}$ )	CL%

<5.0	90	ABLIKIM	13G	BES3	$J/\psi \rightarrow \phi\eta'$
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$\Gamma(\gamma\text{e}^+\text{e}^-)/\Gamma_{\text{total}}$	$\Gamma_{20}/\Gamma$
VALUE (units $10^{-3}$ )	CL%

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.9	90	BRIERE	00	CLEO	10.6 $\text{e}^+\text{e}^-$
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$\Gamma(\gamma\text{e}^+\text{e}^-)/\Gamma(\gamma\gamma)$	$\Gamma_{20}/\Gamma_6$
VALUE (units $10^{-2}$ )	EVTS

2.13±0.09±0.07	864	ABLIKIM	15o	BES3	$J/\psi \rightarrow \gamma\text{e}^+\text{e}^-$
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See key on page 601

# Meson Particle Listings

## $\eta'(958)$

$\Gamma(\pi^0\gamma\gamma)/\Gamma(\pi^0\pi^0\eta)$		$\Gamma_{21}/\Gamma_3$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
<37	90	ALDE 87B GAM2 $38\pi^-p \rightarrow n4\gamma$

$\Gamma(4\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{22}/\Gamma$
VALUE	CL%	DOCUMENT ID TECN COMMENT
<3.2 $\times 10^{-4}$	90	DONSKOV 14 GAM4 $32.5\pi^-p \rightarrow \eta' n$

$\Gamma(4\pi^0)/\Gamma(\pi^0\pi^0\eta)$		$\Gamma_{22}/\Gamma_3$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<23	90	ALDE 87B GAM2 $38\pi^-p \rightarrow n8\gamma$

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$		$\Gamma_{23}/\Gamma$
VALUE	CL%	DOCUMENT ID TECN COMMENT
< 5.6 $\times 10^{-9}$	90	<sup>1</sup> ACHASOV 15 SND $0.958 e^+e^- \rightarrow \pi\pi\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<12 $\times 10^{-9}$	90	<sup>2</sup> AKHMETSHIN 15 CMD3 $0.958 e^+e^- \rightarrow \pi^+\pi^-\eta$
<2.1 $\times 10^{-7}$	90	VOROB'YEV 88 ND $e^+e^- \rightarrow \pi^+\pi^-\eta$
<sup>1</sup> Combining data of ACHASOV 15 and AKHMETSHIN 15 and using $\Gamma(\eta') = 0.198 \pm 0.009$ MeV.		
<sup>2</sup> Using $\Gamma_{\eta'(958)} = 198 \pm 9$ keV, $B(\eta'(958) \rightarrow \pi^+\pi^-\eta) = (42.9 \pm 0.7)\%$ , and $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.20)\%$ .		

$\Gamma(\text{invisible})/\Gamma_{\text{total}}$		$\Gamma_{24}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<9.5	90	<sup>1</sup> NAIK 09 CLEO $J/\psi \rightarrow \gamma\eta'$
<sup>1</sup> Not independent of measured value of $\Gamma_{24}/\Gamma_1$ from NAIK 09.		

$\Gamma(\text{invisible})/\Gamma(\gamma\gamma)$		$\Gamma_{24}/\Gamma_6$
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID TECN COMMENT
<2.4	90	ABLIKIM 13 BES3 $J/\psi \rightarrow \phi\eta'$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<6.69	90	ABLIKIM 06Q BES $J/\psi \rightarrow \phi\eta'$

$\Gamma(\text{invisible})/\Gamma(\pi^+\pi^-\eta)$		$\Gamma_{24}/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<2.1	90	<sup>1</sup> NAIK 09 CLEO $J/\psi \rightarrow \gamma\eta'$
<sup>1</sup> NAIK 09 reports $[\Gamma(\eta'(958) \rightarrow \text{invisible})/\Gamma(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)] < 5.4 \times 10^{-3}$ which we multiply by our best value $B(\eta \rightarrow 2\gamma) = 39.41 \times 10^{-2}$ .		

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{25}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
< 0.6	90	<sup>1</sup> ABLIKIM 11G BES3 $J/\psi \rightarrow \gamma\pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
< 29	90	<sup>2</sup> MORI 07A BELL $\gamma\gamma \rightarrow \pi^+\pi^-$
< 3.3	90	<sup>3</sup> MORI 07A BELL $\gamma\gamma \rightarrow \pi^+\pi^-$
<800	95	DANBURG 73 HBC $2.2 K^-p \rightarrow \Lambda X^0$
<200	90	RITTENBERG 69 HBC $1.7\text{--}2.7 K^-p$
<sup>1</sup> ABLIKIM 11G reports $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta'(958))] < 2.84 \times 10^{-7}$ which we divide by our best value $B(J/\psi(1S) \rightarrow \gamma\eta'(958)) = 5.15 \times 10^{-3}$ .		
<sup>2</sup> Taking into account interference with the $\gamma\gamma \rightarrow \pi^+\pi^-$ continuum.		
<sup>3</sup> Without interference with the $\gamma\gamma \rightarrow \pi^+\pi^-$ continuum.		

$\Gamma(\pi^0\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{26}/\Gamma$
VALUE	CL%	DOCUMENT ID TECN COMMENT
<4 $\times 10^{-4}$	90	<sup>1</sup> ABLIKIM 11G BES3 $J/\psi \rightarrow \gamma\pi^0\pi^0$
<sup>1</sup> ABLIKIM 11G reports $[\Gamma(\eta'(958) \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta'(958))] < 2.84 \times 10^{-7}$ which we divide by our best value $B(J/\psi(1S) \rightarrow \gamma\eta'(958)) = 5.15 \times 10^{-3}$ .		

$\Gamma(\pi^0\pi^0)/\Gamma(\pi^0\pi^0\eta)$		$\Gamma_{26}/\Gamma_3$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
<45	90	ALDE 87B GAM2 $38\pi^-p \rightarrow n4\gamma$

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$		$\Gamma_{27}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID TECN COMMENT
< 1.4	90	BRIERE 00 CLEO $10.6 e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<13	90	RITTENBERG 65 HBC $2.7 K^-p$

$\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$		$\Gamma_{28}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID TECN COMMENT
< 2.4	90	BRIERE 00 CLEO $10.6 e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<11	90	RITTENBERG 65 HBC $2.7 K^-p$

$\Gamma(3\gamma)/\Gamma(\pi^0\pi^0\eta)$		$\Gamma_{29}/\Gamma_3$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
<4.6	90	ALDE 87B GAM2 $38\pi^-p \rightarrow n3\gamma$

$\Gamma(\mu^+\mu^-\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{30}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID TECN COMMENT
<6.0	90	DZHELYADIN 81 CNTR $30\pi^-p \rightarrow \eta' n$

$\Gamma(\mu^+\mu^-\eta)/\Gamma_{\text{total}}$		$\Gamma_{31}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID TECN COMMENT
<1.5	90	DZHELYADIN 81 CNTR $30\pi^-p \rightarrow \eta' n$

$\Gamma(e\mu)/\Gamma_{\text{total}}$		$\Gamma_{32}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID TECN COMMENT
<4.7	90	BRIERE 00 CLEO $10.6 e^+e^-$

### $\eta'(958) \rightarrow \eta\pi\pi$ DECAY PARAMETERS

$$|\text{MATRIX ELEMENT}|^2 = |1 + \alpha Y|^2 + CX + DX^2$$

$X$  and  $Y$  are Dalitz variables;  $\alpha$  is complex and  $C$ , and  $D$  are real-valued. Parameters  $C$  and  $D$  are not necessarily equal to  $c$  and  $d$ , respectively, in the generalized parameterization following this one. May be different for  $\eta'(958) \rightarrow \eta\pi^+\pi^-$  and  $\eta'(958) \rightarrow \eta\pi^0\pi^0$  decays. Because of different initial assumptions and strong correlations of the parameters we do not average the parameters in the section below.

### Re( $\alpha$ ) decay parameter

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.033 \pm 0.005 \pm 0.003$	44k	<sup>1</sup> ABLIKIM 11	BES3	$J/\psi \rightarrow \gamma\eta\pi^+\pi^-$
$-0.072 \pm 0.012 \pm 0.006$	7k	<sup>2</sup> AMELIN 05A	VES	$28\pi^-A \rightarrow \eta\pi^+\pi^-\pi^-A^*$
$-0.021 \pm 0.018 \pm 0.017$	6.7k	<sup>3</sup> BRIERE 00	CLEO	$10.6 e^+e^- \rightarrow \eta\pi^+\pi^-X$
$-0.058 \pm 0.013 \pm 0.003$	5.4k	<sup>4</sup> ALDE 86	GAM2	$38\pi^-p \rightarrow n\eta\pi^0\pi^0$
$-0.08 \pm 0.03$		<sup>4,5</sup> KALBFLEISCH 74	RVUE	$\eta' \rightarrow \eta\pi^+\pi^-$
<sup>1</sup> See ABLIKIM 11 for the full correlation matrix.				
<sup>2</sup> Superseded by DOROFEEV 07, which found this parameterization unacceptable. See below.				
<sup>3</sup> Assuming $\text{Im}(\alpha) = 0$ , $C = 0$ , and $D = 0$ .				
<sup>4</sup> Assuming $C = 0$ .				
<sup>5</sup> From the data of DAUBER 64, RITTENBERG 69, AGUILAR-BENITEZ 72B, JACOBS 73, and DANBURG 73.				

### Im( $\alpha$ ) decay parameter

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.000 \pm 0.049 \pm 0.001$	44k	<sup>1</sup> ABLIKIM 11	BES3	$J/\psi \rightarrow \gamma\eta\pi^+\pi^-$
$0.0 \pm 0.1 \pm 0.0$	7k	<sup>2</sup> AMELIN 05A	VES	$28\pi^-A \rightarrow \eta\pi^+\pi^-\pi^-A^*$
$-0.00 \pm 0.13 \pm 0.00$	5.4k	<sup>3</sup> ALDE 86	GAM2	$38\pi^-p \rightarrow n\eta\pi^0\pi^0$
$0.0 \pm 0.3$		<sup>3,4</sup> KALBFLEISCH 74	RVUE	$\eta' \rightarrow \eta\pi^+\pi^-$
<sup>1</sup> See ABLIKIM 11 for the full correlation matrix.				
<sup>2</sup> Superseded by DOROFEEV 07, which found this parameterization unacceptable. See below.				
<sup>3</sup> Assuming $C = 0$ .				
<sup>4</sup> From the data of DAUBER 64, RITTENBERG 69, AGUILAR-BENITEZ 72B, JACOBS 73, and DANBURG 73.				

### C decay parameter

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$+0.018 \pm 0.009 \pm 0.003$	44k	<sup>1</sup> ABLIKIM 11	BES3	$J/\psi \rightarrow \gamma\eta\pi^+\pi^-$
$0.020 \pm 0.018 \pm 0.004$	7k	<sup>2</sup> AMELIN 05A	VES	$28\pi^-A \rightarrow \eta\pi^+\pi^-\pi^-A^*$
<sup>1</sup> See ABLIKIM 11 for the full correlation matrix.				
<sup>2</sup> Superseded by DOROFEEV 07, which found this parameterization unacceptable. See below.				

### D decay parameter

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.059 \pm 0.012 \pm 0.004$	44k	<sup>1</sup> ABLIKIM 11	BES3	$J/\psi \rightarrow \gamma\eta\pi^+\pi^-$
$-0.066 \pm 0.030 \pm 0.015$	7k	<sup>2</sup> AMELIN 05A	VES	$28\pi^-A \rightarrow \eta\pi^+\pi^-\pi^-A^*$
$0.00 \pm 0.03 \pm 0.00$	5.4k	<sup>3</sup> ALDE 86	GAM2	$38\pi^-p \rightarrow n\eta\pi^0\pi^0$
0		<sup>3,4</sup> KALBFLEISCH 74	RVUE	$\eta' \rightarrow \eta\pi^+\pi^-$
<sup>1</sup> See ABLIKIM 11 for the full correlation matrix.				
<sup>2</sup> Superseded by DOROFEEV 07, which found this parameterization unacceptable. See below.				
<sup>3</sup> Assuming $C = 0$ .				
<sup>4</sup> From the data of DAUBER 64, RITTENBERG 69, AGUILAR-BENITEZ 72B, JACOBS 73, and DANBURG 73.				





# Meson Particle Listings

$f_0(980)$

65 ± 13	262 ± 30	<sup>9</sup> AUBERT	07AK BABR	10.6 $e^+e^- \rightarrow \phi\pi^+\pi^-\gamma$
81 ± 21	54 ± 9	<sup>9</sup> AUBERT	07AK BABR	10.6 $e^+e^- \rightarrow \phi\pi^0\pi^0\gamma$
51.3 <sup>+20.8+13.2</sup> <sub>-17.7-3.8</sub>		<sup>10</sup> MORI	07 BELL	10.6 $e^+e^- \rightarrow \phi\pi^0\pi^0\gamma$
61 ± 9	<sup>+14</sup> <sub>-8</sub>	2584	<sup>11</sup> GARMASH	05 BELL $B^+ \rightarrow K^+\pi^+\pi^-$
64 ± 16		<sup>12</sup> ANISOVICH	03 RVUE	
121 ± 23		TIKHOMIROV	03 SPEC	40.0 $\pi^- \rightarrow K_S^0 K_S^0 \bar{K}_L^0 X$
~ 70		<sup>13</sup> BRAMON	02 RVUE	1.02 $e^+e^- \rightarrow \pi^0\pi^0\gamma$
44 ± 2 ± 2	848	<sup>14</sup> AITALA	01A E791	$D_s^+ \rightarrow \pi^-\pi^+\pi^+$
201 ± 28	419	<sup>15</sup> ACHASOV	00H SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
122 ± 13	419	<sup>16,17</sup> ACHASOV	00H SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
56 ± 20		<sup>18</sup> AKHMETSHIN	99C CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
65 ± 20		BARBERIS	99 OMEG	450 $pp \rightarrow p_S p_f K^+ K^-$
80 ± 10		BARBERIS	99B OMEG	450 $pp \rightarrow p_S p_f \pi^+ \pi^-$
80 ± 10		BARBERIS	99C OMEG	450 $pp \rightarrow p_S p_f \pi^0 \pi^0$
48 ± 12 ± 8		<sup>19</sup> BARBERIS	99D OMEG	450 $pp \rightarrow K^+ K^-, \pi^+ \pi^-$
65 ± 25		BELLAZZINI	99 GAM4	450 $pp \rightarrow p p \pi^0 \pi^0$
71 ± 14		<sup>20</sup> KAMINSKI	99 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, \sigma\sigma$
~ 28		<sup>20</sup> OLLER	99 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
~ 25		OLLER	99B RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
~ 14		OLLER	99C RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, \eta\eta$
70 ± 20		ALDE	98 GAM4	
86 ± 16		<sup>20</sup> ANISOVICH	98B RVUE	Compilation
54		<sup>21</sup> LOCHER	98 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
69 ± 15		<sup>22</sup> ALDE	97 GAM2	450 $pp \rightarrow p p \pi^0 \pi^0$
38 ± 20		<sup>23</sup> BERTIN	97C OBLX	0.0 $\bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
~ 100		<sup>24</sup> ISHIDA	96 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
34		TORNQVIST	96 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, K\pi, \eta\pi$
48 ± 10	3k	<sup>25</sup> ALDE	95B GAM2	38 $\pi^- p \rightarrow \pi^0 \pi^0 n$
95 ± 20	10k	<sup>26</sup> ALDE	95B GAM2	38 $\pi^- p \rightarrow \pi^0 \pi^0 n$
26 ± 10		AMSLER	95B CBAR	0.0 $\bar{p}p \rightarrow 3\pi^0$
~ 112		AMSLER	95D CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta \eta, \pi^0 \pi^0 \eta$
80 ± 12		<sup>28</sup> ANISOVICH	95 RVUE	
70		JANSSEN	95 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
34		<sup>29</sup> BUGG	94 RVUE	$\bar{p}p \rightarrow \eta 2\pi^0$
74		<sup>30</sup> KAMINSKI	94 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
29 ± 2		<sup>31</sup> ZOU	94B RVUE	
46		<sup>32</sup> MORGAN	93 RVUE	$\pi\pi(K\bar{K}) \rightarrow \pi\pi(K\bar{K}), J/\psi \rightarrow \phi\pi\pi(K\bar{K}), D_s \rightarrow \pi(\pi\pi)$
37.4 ± 10.6		<sup>22</sup> AGUILAR-...	91 EHS	400 $pp$
72 ± 8		<sup>33</sup> ARMSTRONG	91 OMEG	300 $pp \rightarrow p p \pi\pi, p p K\bar{K}$
110 ± 30		BREAKSTONE	90 SFM	$pp \rightarrow p p \pi^+ \pi^-$
29 ± 13		<sup>22</sup> ABACHI	86B HRS	$e^+e^- \rightarrow \pi^+ \pi^- X$
120 ± 281 ± 20		ETKIN	82B MPS	23 $\pi^- p \rightarrow n 2K_S^0$
28 ± 10		<sup>33</sup> GIDAL	81 MRK2	$J/\psi \rightarrow \pi^+ \pi^- X$
70 to 300		<sup>34</sup> ACHASOV	80 RVUE	
100 ± 80		<sup>35</sup> AGUILAR-...	78 HBC	0.7 $\bar{p}p \rightarrow K_S^0 K_S^0$
30 ± 8		<sup>33</sup> LEEPER	77 ASPK	2-2.4 $\pi^- p \rightarrow \pi^+ \pi^- n, K^+ K^- n$
48 ± 14		<sup>33</sup> BINNIE	73 CNTR	$\pi^- p \rightarrow n \text{MM}$
32 ± 10		<sup>36</sup> GRAYER	73 ASPK	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
30 ± 10		<sup>36</sup> HYAMS	73 ASPK	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
54 ± 16		<sup>36</sup> PROTOPOP...	73 HBC	7 $\pi^+ p \rightarrow \pi^+ \pi^+ \pi^-$

<sup>1</sup> Analytic continuation using Roy equations. Uses the  $K_{e4}$  data of BATLEY 10C and the  $\pi N \rightarrow \pi\pi N$  data of HYAMS 73, GRAYER 74, and PROTOPODESCU 73.

<sup>2</sup> Quoted number refers to twice imaginary part of pole position.

<sup>3</sup> Analytic continuation using GKPY equations. Uses the  $K_{e4}$  data of BATLEY 10C and the  $\pi N \rightarrow \pi\pi N$  data of HYAMS 73, GRAYER 74, and PROTOPODESCU 73.

<sup>4</sup> Pole position. Used Roy equations.

<sup>5</sup> Average of the analyses of three data sets in the K-matrix model. Uses the data of BATLEY 08A, HYAMS 73, and GRAYER 74, partially of COHEN 80 or ETKIN 82B.

<sup>6</sup> On sheet II in a 2-pole solution. The other pole is found on sheet III at (850–100i) MeV

<sup>7</sup> Using a relativistic Breit-Wigner function and taking into account the finite  $D_S$  mass.

<sup>8</sup> Breit-Wigner  $\pi\pi$  width. Using finite width corrections according to FLATTE 76 and ACHASOV 05, and the ratio  $g_{f_0} K K / g_{f_0} \pi\pi = 0$ .

<sup>9</sup> Systematic errors not estimated.

<sup>10</sup> Breit-Wigner  $\pi\pi$  width. Using finite width corrections according to FLATTE 76 and ACHASOV 05, and the ratio  $g_{f_0} K K / g_{f_0} \pi\pi = 4.21 \pm 0.25 \pm 0.21$  from ABLIKIM 05.

<sup>11</sup> Breit-Wigner, solution 1, PWA ambiguous.

<sup>12</sup> K-matrix pole from combined analysis of  $\pi^- p \rightarrow \pi^0 \pi^0 n$ ,  $\pi^- p \rightarrow \pi^+ \pi^- n$ ,  $\pi^+ \pi^- \rightarrow \pi^0 \pi^0 \pi^0$ ,  $\pi^0 \eta \eta$ ,  $\pi^0 \pi^0 \eta$ ,  $\pi^+ \pi^- \pi^0$ ,  $K^+ K^- \pi^0$ ,  $K_S^0 K_S^0 \pi^0$ ,  $K^+ K_S^0 \pi^-$  at rest,  $\bar{p} n \rightarrow \pi^- \pi^- \pi^+$ ,  $K_S^0 K^- \pi^0$ ,  $K_S^0 K_S^0 \pi^-$  at rest.

<sup>13</sup> Using the data of AKHMETSHIN 99C, ACHASOV 00H, and ALOISIO 02D.

<sup>14</sup> Breit-Wigner width.

<sup>15</sup> Supersedes ACHASOV 98i. Using the model of ACHASOV 89.

<sup>16</sup> Supersedes ACHASOV 98i.

<sup>17</sup> In the “narrow resonance” approximation.

<sup>18</sup> From the combined fit of the photon spectra in the reactions  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma$ ,  $\pi^0 \pi^0 \gamma$ .

<sup>19</sup> Supersedes BARBERIS 99 and BARBERIS 99B

<sup>20</sup> T-matrix pole.

<sup>21</sup> On sheet II in a 2 pole solution. The other pole is found on sheet III at (1039–93i) MeV.

<sup>22</sup> From invariant mass fit.

<sup>23</sup> On sheet II in a 2 pole solution. The other pole is found on sheet III at (963–29i) MeV.

<sup>24</sup> Reanalysis of data from HYAMS 73, GRAYER 74, SRINIVASAN 75, and ROSSELET 77 using the interfering amplitude method.

<sup>25</sup> At high  $|t|$ .

<sup>26</sup> At low  $|t|$ .

<sup>27</sup> On sheet II in a 4-pole solution, the other poles are found on sheet III at (953–55i) MeV and on sheet IV at (938–35i) MeV.

<sup>28</sup> Combined fit of ALDE 95B, ANISOVICH 94,

<sup>29</sup> On sheet II in a 2 pole solution. The other pole is found on sheet III at (996–103i) MeV.

<sup>30</sup> From sheet II pole position.

<sup>31</sup> On sheet II in a 2 pole solution. The other pole is found on sheet III at (797–185i) MeV and can be interpreted as a shadow pole.

<sup>32</sup> On sheet II in a 2 pole solution. The other pole is found on sheet III at (978–28i) MeV.

<sup>33</sup> From coupled channel analysis.

<sup>34</sup> Coupled channel analysis with finite width corrections.

<sup>35</sup> From coupled channel fit to the HYAMS 73 and PROTOPODESCU 73 data. With a simultaneous fit to the  $\pi\pi$  phase-shifts, inelasticity and to the  $K_S^0 K_S^0$  invariant mass.

<sup>36</sup> Included in AGUILAR-BENITEZ 78 fit.

## $f_0(980)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi\pi$	dominant
$\Gamma_2$ $K\bar{K}$	seen
$\Gamma_3$ $\gamma\gamma$	seen
$\Gamma_4$ $e^+e^-$	

## $f_0(980)$ PARTIAL WIDTHS

$\Gamma(\gamma\gamma)$				$\Gamma_3$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
<b>0.31<sup>+0.05</sup><sub>-0.04</sub> OUR AVERAGE</b>				
0.32 ± 0.05	<sup>1</sup> DAI	14A	RVUE	Compilation
0.286 ± 0.017 <sup>+0.211</sup> <sub>-0.070</sub>	<sup>2</sup> UEHARA	08A	BELL	10.6 $e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
0.205 ± 0.095 ± 0.147 <sub>-0.083-0.117</sub>	<sup>3</sup> MORI	07	BELL	10.6 $e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
0.42 ± 0.06 ± 0.18	<sup>4</sup> OEST	90	JADE	$e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.16 ± 0.01	<sup>5</sup> MENNESSIER	11	RVUE	
0.29 ± 0.21 <sup>+0.02</sup> <sub>-0.07</sub>	<sup>6</sup> MOUSSALLAM	11	RVUE	Compilation
0.42	<sup>7,8</sup> PENNINGTON	08	RVUE	Compilation
0.10	<sup>8,9</sup> PENNINGTON	08	RVUE	Compilation
0.28 <sup>+0.09</sup> <sub>-0.13</sub>	<sup>10</sup> BOGLIONE	99	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-, \pi^0\pi^0$
0.29 ± 0.07 ± 0.12	<sup>11,12</sup> BOYER	90	MRK2	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
0.31 ± 0.14 ± 0.09	<sup>11,12</sup> MARSISKE	90	CBAL	$e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
0.63 ± 0.14	<sup>13</sup> MORGAN	90	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-, \pi^0\pi^0$

<sup>1</sup> Using dispersive analysis with phases from GARCIA-MARTIN 11A and BUETTIKER 04 as input.

<sup>2</sup> Using finite width corrections according to FLATTE 76 and ACHASOV 05, and the ratio  $g_{f_0} K K / g_{f_0} \pi\pi = 0$ .

<sup>3</sup> Using finite width corrections according to FLATTE 76 and ACHASOV 05, and the ratio  $g_{f_0} K K / g_{f_0} \pi\pi = 4.21 \pm 0.25 \pm 0.21$  from ABLIKIM 05.

<sup>4</sup> OEST 90 quote systematic errors<sup>+0.08</sup><sub>-0.18</sub>. We use ±0.18. Observed 60 events.

<sup>5</sup> Uses an analytic K-matrix model. Compilation.

<sup>6</sup> Using dispersion integral with phase input from Roy equations and data from MARSISKE 90, BOYER 90, BEHREND 92, UEHARA 08A, and MORI 07.

<sup>7</sup> Solution A (preferred solution based on  $\chi^2$ -analysis).

<sup>8</sup> Dispersion theory based amplitude analysis of BOYER 90, MARSISKE 90, BEHREND 92, and MORI 07.

<sup>9</sup> Solution B (worse than solution A; still acceptable when systematic uncertainties are included).

<sup>10</sup> Supersedes MORGAN 90.

<sup>11</sup> From analysis allowing arbitrary background unconstrained by unitarity.

<sup>12</sup> Data included in MORGAN 90, BOGLIONE 99 analyses.

<sup>13</sup> From amplitude analysis of BOYER 90 and MARSISKE 90, data corresponds to resonance parameters  $m = 989$  MeV,  $\Gamma = 61$  MeV.

$\Gamma(e^+e^-)$				$\Gamma_4$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;8.4</b>	90	VOROBYEV	88	ND $e^+e^- \rightarrow \pi^0\pi^0$



See key on page 601

## Meson Particle Listings

 $f_0(980)$ ,  $a_0(980)$  $f_0(980)$  BRANCHING RATIOS $\Gamma(\pi\pi)/[\Gamma(\pi\pi) + \Gamma(K\bar{K})]$  $\Gamma_1/(\Gamma_1 + \Gamma_2)$ 

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.52 \pm 0.12$	9.9k	<sup>1</sup> AUBERT	06o BABR	$B^{\pm} \rightarrow K^{\pm} \pi^{\pm} \pi^{\mp}$
$0.75^{+0.11}_{-0.13}$		<sup>2</sup> ABLIKIM	05q BES2	$\chi_{c0} \rightarrow 2\pi^{+} 2\pi^{-}$ , $\pi^{+} \pi^{-} K^{+} K^{-}$
$0.84 \pm 0.02$		<sup>3</sup> ANISOVICH	02b SPEC	Combined fit
$\sim 0.68$		OLLER	99b RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}$
$0.67 \pm 0.09$		<sup>4</sup> LOVERRE	80 HBC	$4\pi^{-} p \rightarrow n 2K_S^0$
$0.81^{+0.09}_{-0.04}$		<sup>4</sup> CASON	78 STRC	$7\pi^{-} p \rightarrow n 2K_S^0$
$0.78 \pm 0.03$		<sup>4</sup> WETZEL	76 OSPK	$8.9\pi^{-} p \rightarrow n 2K_S^0$

<sup>1</sup> Recalculated by us using  $\Gamma(K^{+}K^{-}) / \Gamma(\pi^{+}\pi^{-}) = 0.69 \pm 0.32$  from AUBERT 06o and isospin relations.

<sup>2</sup> Using data from ABLIKIM 04g.

<sup>3</sup> From a combined K-matrix analysis of Crystal Barrel ( $0. p\bar{p} \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta\eta, \pi^0 \pi^0 \eta$ ), GAMS ( $\pi p \rightarrow \pi^0 \pi^0 n, \eta\eta n, \eta\eta' n$ ), and BNL ( $\pi p \rightarrow K\bar{K} n$ ) data.

<sup>4</sup> Measure  $\pi\pi$  elasticity assuming two resonances coupled to the  $\pi\pi$  and  $K\bar{K}$  channels only.

MORGAN	90	ZPHY C48 623	D. Morgan, M.R. Pennington	(RAL, DURH)
OEST	90	ZPHY C47 343	T. Oest <i>et al.</i>	(JADE Collab.)
ACHASOV	89	NP B315 465	N.N. Achasov, V.N. Ivanchenko	
AUGUSTIN	89	NP B320 1	J.E. Augustin, G. Cosme	(DM2 Collab.)
VOROBYEV	88	SJNP 48 273	P.V. Vorobiev <i>et al.</i>	(NOVO)
		Translated from YAF 48 436.		
ABACHI	86b	PRL 57 1930	S. Abachi <i>et al.</i>	(PURD, ANL, IND, MICH+)
ETKIN	82b	PR D25 1786	A. Etkin <i>et al.</i>	(BNL, CUNY, TUFTS, VAND)
GIDAL	81	PL 107B 153	G. Gidal <i>et al.</i>	(SLAC, LBL)
ACHASOV	80	SJNP 32 566	N.N. Achasov, S.A. Devyanin, G.N. Shestakov	(NOVM)
		Translated from YAF 32 109b.		
COHEN	80	PR D22 2595	D. Cohen <i>et al.</i>	(ANL) IJP
LOVERRE	80	ZPHY C6 187	P.F. Loverre <i>et al.</i>	(CERN, CDEF, MADR+) IJP
AGUILAR...	78	NP B140 73	M. Aguilar-Benitez <i>et al.</i>	(MADR, BOMB+)
CASON	78	PRL 41 271	N.M. Cason <i>et al.</i>	(NDAM, ANL)
LEEPER	77	PR D16 2054	R.J. Leeper <i>et al.</i>	(ISU)
ROSSELET	77	PR D15 574	L. Rosselet <i>et al.</i>	(GEVA, SACL)
FLATTE	76	PL 63B 224	S.M. Flatte	(CERN)
WETZEL	76	NP B115 208	W. Wetzel <i>et al.</i>	(ETH, CERN, LOIC)
SRINIVASAN	75	PR D12 681	V. Srinivasan <i>et al.</i>	(NDAM, ANL)
GRAY	74	NP B75 189	G. Grayer <i>et al.</i>	(CERN, MPIM)
BINNIE	73	PRL 31 1534	D.M. Binnie <i>et al.</i>	(LOIC, SHMP)
GRAY	73	Tallahassee	G. Grayer <i>et al.</i>	(CERN, MPIM)
HYAMS	73	NP B64 134	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
PROTOPOP...	73	PR D7 1279	S.D. Protopopescu <i>et al.</i>	(LBL)

 $a_0(980)$ 

$$I^G(J^{PC}) = 1^-(0^{+} +)$$

See our minireview on scalar mesons under  $f_0(500)$ . (See the index for the page number.)

 $a_0(980)$  MASS

VALUE (MeV)	DOCUMENT ID
<b>980 ± 20 OUR ESTIMATE</b>	Mass determination very model dependent

 $\eta\pi$  FINAL STATE ONLY

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$982.5 \pm 1.6 \pm 1.1$	16.9k	<sup>1</sup> AMBROSINO	09f KLOE		$1.02 e^{+} e^{-} \rightarrow \eta\pi^0 \gamma$
$986 \pm 4$		ANISOVICH	09 RVUE		$0.0 \bar{p}p, \pi N$
$982.3 \pm 0.6^{+3.1}_{-0.7}$		<sup>2</sup> UEHARA	09a BELL		$\gamma\gamma \rightarrow \pi^0 \eta$
$987.4 \pm 1.0 \pm 3.0$		<sup>3,4</sup> BUGG	08a RVUE 0		$\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
$989.1 \pm 1.0 \pm 3.0$		<sup>4,5</sup> BUGG	08a RVUE 0		$\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
$985 \pm 4 \pm 6$	318	ACHARD	02b L3		$183-209 e^{+} e^{-} \rightarrow e^{+} e^{-} \eta\pi^{+} \pi^{-}$
$995^{+52}_{-10}$	36	<sup>6</sup> ACHASOV	00f SND		$e^{+} e^{-} \rightarrow \eta\pi^0 \gamma$
$994^{+33}_{-8}$	36	<sup>7</sup> ACHASOV	00f SND		$e^{+} e^{-} \rightarrow \eta\pi^0 \gamma$
$975 \pm 7$		BARBERIS	00h		$450 pp \rightarrow p\eta\pi^0 p_S$
$988 \pm 8$		BARBERIS	00h		$450 pp \rightarrow \Delta^{++} \eta\pi^{-} p_S$
$\sim 1055$		<sup>8</sup> OLLER	99 RVUE		$\eta\pi, K\bar{K}$
$\sim 1009.2$		<sup>8</sup> OLLER	99b RVUE		$\pi\pi \rightarrow \pi\pi, K\bar{K}$
$993.1 \pm 2.1$		<sup>9</sup> TEIGE	99 B852		$18.3 \pi^{-} p \rightarrow \eta\pi^{+} \pi^{-} n$
$988 \pm 6$		<sup>8</sup> ANISOVICH	98b RVUE		Compilation
$987$		TORNQVIST	96 RVUE		$\pi\pi \rightarrow \pi\pi, K\bar{K}, K\pi, \eta\pi$
$991$		JANSSEN	95 RVUE		$\eta\pi \rightarrow \eta\pi, K\bar{K}, K\pi, \eta\pi$
$984.45 \pm 1.23 \pm 0.34$		<sup>10</sup> AMSLER	94c CBAR		$0.0 \bar{p}p \rightarrow \omega\eta\pi^0$
$982 \pm 2$		<sup>10</sup> AMSLER	92 CBAR		$0.0 \bar{p}p \rightarrow \eta\eta\pi^0$
$984 \pm 4$	1040	<sup>10</sup> ARMSTRONG	91b OMEG ±		$300 pp \rightarrow p\eta\pi^{+} \pi^{-}$
$976 \pm 6$		ATKINSON	84e OMEG ±		$25-55 \gamma p \rightarrow \eta\pi n$
$986 \pm 3$	500	<sup>11</sup> EVANGELIS...	81 OMEG ±		$12 \pi^{-} p \rightarrow \eta\pi^{+} \pi^{-} \pi^{-} p$
$990 \pm 7$	145	<sup>11</sup> GURTU	79 HBC ±		$4.2 K^{-} p \rightarrow \Lambda\eta 2\pi$
$980 \pm 11$	47	CONFORTO	78 OSPK ±		$4.5 \pi^{-} p \rightarrow pX^{-}$
$978 \pm 16$	50	CORDEN	78 OMEG ±		$12-15 \pi^{-} p \rightarrow n\eta 2\pi$
$977 \pm 7$		GRASSLER	77 HBC ±		$16 \pi^{\mp} p \rightarrow p\eta 3\pi$
$989 \pm 4$	70	WELLS	75 HBC ±		$3.1-6 K^{-} p \rightarrow \Lambda\eta 2\pi$
$972 \pm 10$	150	DEFOIX	72 HBC ±		$0.7 \bar{p}p \rightarrow 7\pi$
$970 \pm 15$	20	BARNES	69c HBC ±		$4-5 K^{-} p \rightarrow \Lambda\eta 2\pi$
$980 \pm 10$		CAMPBELL	69 DBC ±		$2.7 \pi^{+} d$
$980 \pm 10$	15	MILLER	69b HBC ±		$4.5 K^{-} N \rightarrow \eta\pi\Lambda$
$980 \pm 10$	30	AMMAR	68 HBC ±		$5.5 K^{-} p \rightarrow \Lambda\eta 2\pi$

<sup>1</sup> Using the model of ACHASOV 89 and ACHASOV 03b.

<sup>2</sup> From a fit with the S-wave amplitude including two interfering Breit-Wigners plus a background term.

<sup>3</sup> Parameterizes couplings to  $\bar{K}K, \pi\eta$ , and  $\pi\eta'$ .

<sup>4</sup> Using AMSLER 94d and ABELE 98.

<sup>5</sup> From the T-matrix pole on sheet II.

<sup>6</sup> Using the model of ACHASOV 89. Supersedes ACHASOV 98b.

<sup>7</sup> Using the model of JAFFE 77. Supersedes ACHASOV 98b.

<sup>8</sup> T-matrix pole.

<sup>9</sup> Breit-Wigner fit, average between  $a_0^{\pm}$  and  $a_0^0$ . The fit favors a slightly heavier  $a_0^{\pm}$ .

<sup>10</sup> From a single Breit-Wigner fit.

<sup>11</sup> From  $f_1(1285)$  decay.

 $f_0(980)$  REFERENCES

ABLIKIM	15P	PR D92 012007	M. Ablikim <i>et al.</i>	(BES III Collab.)
DAI	14A	PR D90 036004	L.-Y. Dai, M.R. Pennington	(CEBAF)
ABLIKIM	12E	PRL 108 182001	M. Ablikim <i>et al.</i>	(BES III Collab.)
GARCIA-MAR...	11	PRL 107 072001	R. Garcia-Martin <i>et al.</i>	(MADR, CRAC)
GARCIA-MAR...	11A	PR D83 074004	R. Garcia-Martin <i>et al.</i>	(MADR, CRAC)
MENNESSIER	11	PL B696 40	G. Mennessier, S. Narison, X.-G. Wang	
MOUSSALLAM	11	EPJ C71 1814	B. Moussallam	
BATLEY	10C	EPJ C70 635	J.R. Batley <i>et al.</i>	(CERN NA48/2 Collab.)
MENNESSIER	10	PL B688 59	G. Mennessier, S. Narison, X.-G. Wang	
ANISOVICH	09	UIMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
ECKLUND	09	PR D80 052009	K.M. Ecklund <i>et al.</i>	(CLEO Collab.)
BATLEY	08A	EPJ C54 411	J.R. Batley <i>et al.</i>	(CERN NA48/2 Collab.)
PENNINGTON	08	EPJ C56 1	M.R. Pennington <i>et al.</i>	
UEHARA	08A	PR D78 052004	S. Uehara <i>et al.</i>	(BELLE Collab.)
AMBROSINO	07	EPJ C49 473	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AUBERT	07AK	PR D76 012008	B. Aubert <i>et al.</i>	(BABAR Collab.)
BONVICINI	07	PR D76 012001	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
MORI	07	PR D75 051101	T. Mori <i>et al.</i>	(BELLE Collab.)
AMBROSINO	06B	PL B634 148	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AUBERT	06O	PR D74 032003	B. Aubert <i>et al.</i>	(BABAR Collab.)
GARMASH	06	PRL 96 251803	A. Garmash <i>et al.</i>	(BELLE Collab.)
ABLIKIM	05P	PL B607 243	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05Q	PR D72 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
ACHASOV	05	PR D72 013006	N.N. Achasov, G.N. Shestakov	
GARMASH	05	PR D71 092003	A. Garmash <i>et al.</i>	(BELLE Collab.)
ABLIKIM	04G	PR D70 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
BUETTIGER	04	EPJ C33 409	P. Buettiker, S. Descotes-Genon, B. Moussallam	
ANISOVICH	03	EPJ A16 229	V.V. Anisovich <i>et al.</i>	
TIKHOMIROV	03	PAN 66 828	G.D. Tikhomirov <i>et al.</i>	
		Translated from YAF 66 850.		
ALOISIO	02D	PL B537 21	A. Aloisio <i>et al.</i>	(KLOE Collab.)
ANISOVICH	02D	PAN 65 1545	V.V. Anisovich <i>et al.</i>	
		Translated from YAF 65 1583.		
BRAMON	02	EPJ C26 253	A. Bramon <i>et al.</i>	
ACHASOV	01F	PR D63 094007	N.N. Achasov, V.V. Gubin	(Novosibirsk SND Collab.)
AITALA	01A	PRL 86 765	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	01B	PRL 86 770	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AKHMETSHIN	00H	PL B485 349	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	99B	PL B462 371	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99C	PL B462 380	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BARBERIS	99P	PL B453 305	D. Barberis <i>et al.</i>	(Omega Expt.)
BARBERIS	99B	PL B453 316	D. Barberis <i>et al.</i>	(Omega Expt.)
BARBERIS	99C	PL B453 325	D. Barberis <i>et al.</i>	(Omega Expt.)
BARBERIS	99D	PL B462 462	D. Barberis <i>et al.</i>	(Omega Expt.)
BELLAZZINI	99	PL B467 296	R. Bellazzini <i>et al.</i>	
BOGLIONE	99	EPJ C9 11	M. Boglione, M.R. Pennington	
KAMINSKI	99	EPJ C9 141	R. Kaminski, L. Lesniak, B. Loiseau	(CRAC, PARIN)
OLLER	99	PR D60 099906 (erratum)	J.A. Oller <i>et al.</i>	
OLLER	99B	NP A652 407 (erratum)	J.A. Oller, E. Oset	
OLLER	99C	PR D60 074023	J.A. Oller, E. Oset	
ACHASOV	98I	PL B440 442	M.N. Achasov <i>et al.</i>	
ACKERSTAFF	98Q	EPJ C4 19	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ALDE	98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
		PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
		Translated from YAF 62 446.		
ANISOVICH	98B	SPU 41 419	V.V. Anisovich <i>et al.</i>	
		Translated from UFN 168 481.		
LOCHER	98	EPJ C4 317	M.P. Locher <i>et al.</i>	(PSI)
ALDE	97	PL B397 350	D.M. Alde <i>et al.</i>	(GAMS Collab.)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ISHIDA	96	PTP 95 745	S. Ishida <i>et al.</i>	(TOKYU, MIYA, KEK)
TORNQVIST	96	PRL 76 1575	N.A. Tornqvist, M. Roos	(HELS)
ALDE	95B	ZPHY C66 375	D.M. Alde <i>et al.</i>	(GAMS Collab.)
AMSLER	95B	PL B342 433	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	95D	PL B355 425	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
ANISOVICH	95	PL B355 363	V.V. Anisovich <i>et al.</i>	(PNPI, SERP)
JANSSEN	95	PR D52 2690	G. Janssen <i>et al.</i>	(STON, ADL2, JULI)
AMSLER	94D	PL B333 277	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
ANISOVICH	94	PL B323 233	V.V. Anisovich <i>et al.</i>	(Crystal Barrel Collab.)
BUGG	94	PR D50 4412	D.V. Bugg <i>et al.</i>	(LOQM)
KAMINSKI	94	PR D50 3145	R. Kaminski, L. Lesniak, J.P. Maillet	(CRAC+)
ZOU	94B	EPJ C4 591	B.S. Zou, D.V. Bugg	(LOQM)
MORGAN	93	PR D48 1185	D. Morgan, M.R. Pennington	(RAL, DURH)
BEHREND	92	ZPHY C56 381	H.J. Behrend	(CELLO Collab.)
AGUILAR...	91	ZPHY C50 405	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
ARMSTRONG	91	ZPHY C51 351	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
BOYER	90	PR D42 1350	J. Boyer <i>et al.</i>	(Mark II Collab.)
BREAKSTONE	90	ZPHY C48 569	A.M. Breakstone <i>et al.</i>	(ISU, BNGA, CERN+)
MARSISKE	90	PR D41 3324	H. Marsiske <i>et al.</i>	(Crystal Ball Collab.)

# Meson Particle Listings

## $a_0(980)$

### $K\overline{K}$ ONLY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
~ 1053		<sup>12</sup> OLLER 99c	RVUE		$\pi\pi \rightarrow \pi\pi, K\overline{K}$
982 ± 3		<sup>13</sup> ABELE 98	CBAR		$0.0 \overline{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
975 ± 15		BERTIN 98b	OBLX ±		$0.0 \overline{p}p \rightarrow K^\pm K_S \pi^\mp$
976 ± 6	316	DEBILLY 80	HBC ±		$1.2\text{--}2 \overline{p}p \rightarrow f_1(1285)\omega$
1016 ± 10	100	<sup>14</sup> ASTIER 67	HBC ±		$0.0 \overline{p}p$
1003.3 ± 7.0	143	<sup>15</sup> ROSENFELD 65	RVUE ±		
<sup>12</sup> T-matrix pole. <sup>13</sup> T-matrix pole on sheet II, the pole on sheet III is at 1006-i49 MeV. <sup>14</sup> ASTIER 67 includes data of BARLOW 67, CONFORTO 67, ARMENTEROS 65. <sup>15</sup> Plus systematic errors.					

### $a_0(980)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>50 ± 10 OUR ESTIMATE</b> Width determination very model dependent. Peak width in $\eta\pi$ is about 60 MeV, but decay width can be much larger.					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
75.6 ± 1.6		<sup>16</sup> UEHARA 09A	BELL		$\gamma\gamma \rightarrow \pi^0\eta$
80.2 ± 3.8 ± 5.4		<sup>17</sup> BUGG 08A	RVUE 0		$\overline{p}p \rightarrow \pi^0\pi^0\eta$
50 ± 13 ± 4	318	ACHARD 02b	L3		$183\text{--}209 e^+e^- \rightarrow e^+e^-\eta\pi^+\pi^-$
72 ± 16		BARBERIS 00H			$450 \rho\rho \rightarrow \rho_f\eta\pi^0\rho_S$
61 ± 19		BARBERIS 00H			$450 \rho\rho \rightarrow \Delta_T^+ \eta\pi^-\rho_S$
~ 42		<sup>18</sup> OLLER 99	RVUE		$\eta\pi, K\overline{K}$
~ 112		<sup>18</sup> OLLER 99b	RVUE		$\pi\pi \rightarrow \eta\pi, K\overline{K}$
71 ± 7		TEIGE 99	B852		$18.3 \pi^-\rho \rightarrow \eta\pi^+\pi^-\eta$
92 ± 20		<sup>18</sup> ANISOVICH 98b	RVUE		Compilation
65 ± 10		<sup>19</sup> BERTIN 98b	OBLX ±		$0.0 \overline{p}p \rightarrow K^\pm K_S \pi^\mp$
~ 100		TORNQVIST 96	RVUE		$\pi\pi \rightarrow \pi\pi, K\overline{K}, K\pi, \eta\pi$
202		JANSSEN 95	RVUE		$\eta\pi \rightarrow \eta\pi, K\overline{K}, K\pi, \eta\pi$
54.12 ± 0.34 ± 0.12		AMSLER 94c	CBAR		$0.0 \overline{p}p \rightarrow \omega\eta\pi^0$
54 ± 10		<sup>20</sup> AMSLER 92	CBAR		$0.0 \overline{p}p \rightarrow \eta\eta\pi^0$
95 ± 14	1040	<sup>20</sup> ARMSTRONG 91b	OMEG ±		$300 \rho\rho \rightarrow \rho\rho\eta\pi^+\pi^-$
62 ± 15	500	<sup>21</sup> EVANGELIS... 81	OMEG ±		$12 \pi^-\rho \rightarrow \eta\pi^+\pi^-\pi^-\rho$
60 ± 20	145	<sup>21</sup> GURTU 79	HBC ±		$4.2 K^-\rho \rightarrow \Lambda\eta 2\pi$
60 <sup>+50</sup> <sub>-30</sub>	47	CONFORTO 78	OSPK -		$4.5 \pi^-\rho \rightarrow \rho X^-$
86.0 <sup>+60.0</sup> <sub>-50.0</sub>	50	CORDEN 78	OMEG ±		$12\text{--}15 \pi^-\rho \rightarrow n\eta 2\pi$
44 ± 22		GRASSLER 77	HBC -		$16 \pi^\mp\rho \rightarrow \rho\eta 3\pi$
80 to 300	<sup>22</sup> FLATTE	76	RVUE -		$4.2 K^-\rho \rightarrow \Lambda\eta 2\pi$
16.0 <sup>+25.0</sup> <sub>-16.0</sub>	70	WELLS 75	HBC -		$3.1\text{--}6 K^-\rho \rightarrow \Lambda\eta 2\pi$
30 ± 5	150	DEFOIX 72	HBC ±		$0.7 \overline{p}p \rightarrow 7\pi$
40 ± 15		CAMPBELL 69	DBC ±		$2.7 \pi^+d$
60 ± 30	15	MILLER 69b	HBC -		$4.5 K^-N \rightarrow \eta\pi\Lambda$
80 ± 30	30	AMMAR 68	HBC ±		$5.5 K^-\rho \rightarrow \Lambda\eta 2\pi$

- <sup>16</sup> From a fit with the S-wave amplitude including two interfering Breit-Wigners plus a background term.  
<sup>17</sup> From the T-matrix pole on sheet II, using AMSLER 94d and ABELE 98.  
<sup>18</sup> T-matrix pole.  
<sup>19</sup> The  $\eta\pi$  width.  
<sup>20</sup> From a single Breit-Wigner fit.  
<sup>21</sup> From  $f_1(1285)$  decay.  
<sup>22</sup> Using a two-channel resonance parametrization of GAY 76b data.

### $K\overline{K}$ ONLY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>92 ± 8</b>		<sup>23</sup> ABELE 98	CBAR		$0.0 \overline{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
~ 24		<sup>24</sup> OLLER 99c	RVUE		$\pi\pi \rightarrow \pi\pi, K\overline{K}$
~ 25	100	<sup>25</sup> ASTIER 67	HBC ±		
57 ± 13	143	<sup>26</sup> ROSENFELD 65	RVUE ±		
<sup>23</sup> T-matrix pole on sheet II, the pole on sheet III is at 1006-i49 MeV. <sup>24</sup> T-matrix pole. <sup>25</sup> ASTIER 67 includes data of BARLOW 67, CONFORTO 67, ARMENTEROS 65. <sup>26</sup> Plus systematic errors.					

### $a_0(980)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \eta\pi$	dominant
$\Gamma_2 K\overline{K}$	seen
$\Gamma_3 \rho\pi$	
$\Gamma_4 \gamma\gamma$	seen
$\Gamma_5 e^+e^-$	

### $a_0(980)$ PARTIAL WIDTHS

#### $\Gamma(\gamma\gamma)$

VALUE (keV)	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		
0.30 ± 0.10	<sup>27</sup> AMSLER 98	RVUE
<sup>27</sup> Using $\Gamma_{\gamma\gamma}B(a_0(980) \rightarrow \eta\pi) = 0.24 \pm 0.08$ keV.		

### $a_0(980) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

#### $\Gamma(\eta\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21 <sup>+0.08</sup><sub>-0.04</sub> OUR AVERAGE</b>				
0.128 <sup>+0.003+0.502</sup> <sub>-0.002-0.043</sub>		<sup>28</sup> UEHARA 09A	BELL	$\gamma\gamma \rightarrow \pi^0\eta$
0.28 ± 0.04 ± 0.10	44	OEST 90	JADE	$e^+e^- \rightarrow e^+e^-\pi^0\eta$
0.19 ± 0.07 <sup>+0.10</sup> <sub>-0.07</sub>		ANTREAS YAN 86	CBAL	$e^+e^- \rightarrow e^+e^-\pi^0\eta$
<sup>28</sup> From a fit with the S-wave amplitude including two interfering Breit-Wigners plus a background term.				

#### $\Gamma(\eta\pi) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.5</b>	90	VOROBYEV 88	ND	$e^+e^- \rightarrow \pi^0\eta$

### $a_0(980)$ BRANCHING RATIOS

#### $\Gamma(K\overline{K})/\Gamma(\eta\pi)$

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.183 ± 0.024 OUR AVERAGE</b>	Error includes scale factor of 1.2.			
0.57 ± 0.16	<sup>29</sup> BARGIOTTI 03	OBLX		$\overline{p}p$
0.23 ± 0.05	<sup>30</sup> ABELE 98	CBAR		$0.0 \overline{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
0.166 ± 0.01 ± 0.02	<sup>31</sup> BARBERIS 98c	OMEG		$450 \rho\rho \rightarrow \rho_f f_1(1285) \rho_S$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.20 ± 0.15	<sup>32</sup> ANISOVICH 09	RVUE		$0.0 \overline{p}p, \pi N$
1.05 ± 0.07 ± 0.05	<sup>33</sup> BUGG 08A	RVUE 0		$\overline{p}p \rightarrow \pi^0\pi^0\eta$
~ 0.60	OLLER 99b	RVUE		$\pi\pi \rightarrow \eta\pi, K\overline{K}$
0.7 ± 0.3	<sup>31</sup> CORDEN 78	OMEG		$12\text{--}15 \pi^-\rho \rightarrow n\eta 2\pi$
0.25 ± 0.08	<sup>31</sup> DEFOIX 72	HBC ±		$0.7 \overline{p} \rightarrow 7\pi$

#### $\Gamma(\rho\pi)/\Gamma(\eta\pi)$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.25	70	AMMAR 70	HBC ±		$4.1, 5.5 K^-\rho \rightarrow \Lambda\eta 2\pi$
<sup>29</sup> Coupled channel analysis of $\pi^+\pi^-\pi^0, K^+K^-\pi^0$ , and $K^\pm K_S^0 \pi^\mp$ . <sup>30</sup> Using $\pi^0\pi^0\eta$ from AMSLER 94d. <sup>31</sup> From the decay of $f_1(1285)$ . <sup>32</sup> This is a ratio of couplings. <sup>33</sup> A ratio of couplings, using AMSLER 94d and ABELE 98. Supersedes BUGG 94.					

### $a_0(980)$ REFERENCES

AMBROSINO 09F	PL B681 5	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
ANISOVICH 09	IJMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
UEHARA 09A	PR D80 032001	S. Uehara <i>et al.</i>	(BELLE Collab.)
BUGG 08A	PR D78 074023	D.V. Bugg	(LOQM)
ACHASOV 03B	PR D68 014006	N.N. Achasov, A.V. Kiselev	
BARGIOTTI 03	EPJ C26 371	M. Bargiotti <i>et al.</i>	(OBELIX Collab.)
ACHARD 02B	PL B526 269	P. Achard <i>et al.</i>	(L3 Collab.)
ACHASOV 00F	PL B479 53	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
BARBERIS 00H	PL B488 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
OLLER 99	PR D60 099906 (erratum)	J.A. Oller <i>et al.</i>	
OLLER 99b	NP A652 407 (erratum)	J.A. Oller, E. Oset	
OLLER 99c	PR D60 074023	J.A. Oller, E. Oset	
TEIGE 99	PR D59 012001	S. Teige <i>et al.</i>	(BNL E852 Collab.)
ABELE 98	PR D57 3860	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV 98B	PL B438 441	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AMSLER 98	RMP 70 1293	C. Amsler	
ANISOVICH 98B	SPU 41 419	V.V. Anisovich <i>et al.</i>	
Translated from UFN 168 481.			
BARBERIS 98C	PL B440 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BERTIN 98B	PL B434 180	A. Bertin <i>et al.</i>	(OBELIX Collab.)
TORNQVIST 96	PRL 76 1575	N.A. Tornqvist, M. Roos	(HELS)
JANSSEN 95	PR D52 2690	G. Janssen <i>et al.</i>	(STON, ADDL, JULI)
AMSLER 94C	PL B327 425	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER 94D	PL B333 277	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
BUGG 94	PR D50 4412	D.V. Bugg <i>et al.</i>	(LOQM)
AMSLER 92	PL B291 347	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
ARMSTRONG 91B	ZPHY C52 389	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
OEST 90	ZPHY C47 343	T. Oest <i>et al.</i>	(JADE Collab.)
ACHASOV 89	NP B315 465	N.N. Achasov, V.N. Ivanchenko	
VOROBYEV 88	SJNP 48 273	P.V. Vorobiev <i>et al.</i>	(NOVO)
Translated from YAF 48 436.			
ANTREAS YAN 86	PR D33 1847	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
ATKINSON 84E	PL 138B 459	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
EVANGELIS... 81	NP B178 197	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
DEBILLY 80	NP B176 1	L. de Billy <i>et al.</i>	(CURIN, LAUS, NEUC+)
GURTU 79	NP B151 181	A. Gurtu <i>et al.</i>	(CERN, ZEEM, NIJM, OXF)
CONFORTO 78	LNC 23 419	B. Conforto <i>et al.</i>	(RHEL, TATO, CHC+)
CORDEN 78	NP B144 253	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TEL+)
GRASSLER 77	NP B121 189	H. Grassler <i>et al.</i>	(AACH3, BERL, BONN+)

See key on page 601

## Meson Particle Listings

 $a_0(980)$ ,  $\phi(1020)$ 

JAFFE	77	PR D15 267,281	R. Jaffe	(MIT)
FLATTE	76	PL 63B 224	S.M. Flatte	(CERN)
GAY	76B	PL 63B 220	J.B. Gay <i>et al.</i>	(CERN, AMST, NUM) JP
WELLS	75	NP B101 333	J. Wells <i>et al.</i>	(OXF)
DEFOIX	72	NP B44 125	C. Defoix <i>et al.</i>	(CDEF, CERN)
AMMAR	70	PR D2 430	R. Ammar <i>et al.</i>	(KANS, NWES, ANL, WISC)
BARNES	69C	PRL 23 610	V.E. Barnes <i>et al.</i>	(BNL, SYRA)
CAMPBELL	69	PRL 22 1204	J.H. Campbell <i>et al.</i>	(PURD)
MILLER	69B	PL 29B 255	D.H. Miller <i>et al.</i>	(PURD)
Also		PR 188 2011	W.L. Yen <i>et al.</i>	(PURD)
AMMAR	68	PRL 21 1832	R. Ammar <i>et al.</i>	(NWES, ANL)
ASTIER	67	PL 25B 294	A. Astier <i>et al.</i>	(CDEF, CERN, IRAD)
Includes data of BARLOW 67, CONFORTO 67, and ARMENTEROS 65.				
BARLOW	67	NC 50A 701	J. Barlow <i>et al.</i>	(CERN, CDEF, IRAD, LIVP)
CONFORTO	67	NP B3 469	G. Conforto <i>et al.</i>	(CERN, CDEF, IPNP+)
ARMENTEROS	65	PL 17 344	R. Armenteros <i>et al.</i>	(CERN, CDEF)
ROSENFELD	65	Oxford Conf. 58	A.H. Rosenfeld	(LRL)

 $\phi(1020)$ 

$$I^G(J^{PC}) = 0^-(1^{--})$$

 $\phi(1020)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1019.461 ± 0.019 OUR AVERAGE</b>		Error includes scale factor of 1.1.		
1019.51 ± 0.02 ± 0.05		<sup>1</sup> LEES	13Q	BABR $e^+e^- \rightarrow K^+K^-\gamma$
1019.30 ± 0.02 ± 0.10	105k	AKHMETSHIN 06	CMD2	0.98–1.06 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.52 ± 0.05 ± 0.05	17.4k	AKHMETSHIN 05	CMD2	0.60–1.38 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.483 ± 0.011 ± 0.025	272k	<sup>2</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow K_L^0 K_S^0$
1019.42 ± 0.05	1900k	<sup>3</sup> ACHASOV	01E	SND $e^+e^- \rightarrow K^+K^-, K_S K_L, \pi^+\pi^-\pi^0$
1019.40 ± 0.04 ± 0.05	23k	AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
1019.36 ± 0.12		<sup>4</sup> ACHASOV	00B	SND $e^+e^- \rightarrow \eta\gamma$
1019.38 ± 0.07 ± 0.08	2200	<sup>5</sup> AKHMETSHIN 99F	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.51 ± 0.07 ± 0.10	11169	AKHMETSHIN 98	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
1019.5 ± 0.4		BARBERIS	98	OMEG 450 $pp \rightarrow \pi^+\pi^-\pi^0$
1019.42 ± 0.06	55600	AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow \text{hadrons}$
1019.7 ± 0.3	2012	DAVENPORT	86	MPSF 400 $pA \rightarrow 4KX$
1019.7 ± 0.1 ± 0.1	5079	ALBRECHT	85D	ARG 10 $e^+e^- \rightarrow K^+K^-\pi^0$
1019.3 ± 0.1	1500	ARENTON	82	AEMS 11.8 polar. $pp \rightarrow K K$
1019.67 ± 0.17	25080	<sup>6</sup> PELLINEN	82	RVUE
1019.52 ± 0.13	3681	BUKIN	78C	OLYA $e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1019.48 ± 0.01		LEES	13F	BABR $D^+ \rightarrow K^+K^-\pi^+$
1019.441 ± 0.008 ± 0.080	542k	<sup>7</sup> AKHMETSHIN 08	CMD2	1.02 $e^+e^- \rightarrow K^+K^-\pi^0$
1019.63 ± 0.07	12540	<sup>8</sup> AUBERT, B	05J	BABR $D^0 \rightarrow \bar{K}^0 K^+ K^-$
1019.8 ± 0.7		ARMSTRONG	86	OMEG 85 $\pi^+ / p p \rightarrow \pi^+ / p 4K p$
1020.1 ± 0.11	5526	<sup>8</sup> ATKINSON	86	OMEG 20–70 $\gamma p$
1019.7 ± 1.0		BEBEK	86	CLEO $e^+e^- \rightarrow \gamma(4S)$
1019.411 ± 0.008	642k	<sup>9</sup> DIJKSTRA	86	SPEC 100–200 $\pi^\pm, \bar{p}, p, K^\pm$ , on Be
1020.9 ± 0.2		<sup>8</sup> FRAME	86	OMEG 13 $K^+p \rightarrow \phi K^+ p$
1021.0 ± 0.2		<sup>8</sup> ARMSTRONG	83B	OMEG 18.5 $K^-p \rightarrow K^-K^+\Lambda$
1020.0 ± 0.5		<sup>8</sup> ARMSTRONG	83B	OMEG 18.5 $K^-p \rightarrow K^-K^+\Lambda$
1019.7 ± 0.3		<sup>8</sup> BARATE	83	GOLI 190 $\pi^- \text{Be} \rightarrow 2\mu X$
1019.8 ± 0.2 ± 0.5	766	IVANOV	81	OLYA 1–1.4 $e^+e^- \rightarrow K^+K^-\pi^0$
1019.4 ± 0.5	337	COOPER	78B	HBC 0.7–0.8 $\bar{p} p \rightarrow K_S^0 K_L^0 \pi^+\pi^-$
1020 ± 1	383	<sup>8</sup> BALDI	77	CNTR 10 $\pi^-p \rightarrow \pi^-\phi p$
1018.9 ± 0.6	800	COHEN	77	ASPK 6 $\pi^\pm N \rightarrow K^+K^-N$
1019.7 ± 0.5	454	KALBFLEISCH	76	HBC 2.18 $K^-p \rightarrow \Lambda K \bar{K}$
1019.4 ± 0.8	984	BESCH	74	CNTR 2 $\gamma p \rightarrow p K^+ K^-$
1020.3 ± 0.4	100	BALLAM	73	HBC 2.8–9.3 $\gamma p$
1019.4 ± 0.7		BINNIE	73B	CNTR $\pi^-p \rightarrow \phi n$
1019.6 ± 0.5	120	<sup>10</sup> AGUILAR-...	72B	HBC 3.9, 4.6 $K^-p \rightarrow \Lambda K^+ K^-$
1019.9 ± 0.5	100	<sup>10</sup> AGUILAR-...	72B	HBC 3.9, 4.6 $K^-p \rightarrow K^-p K^+ K^-$
1020.4 ± 0.5	131	COLLEY	72	HBC 10 $K^+p \rightarrow K^+p \phi$
1019.9 ± 0.3	410	STOTTLE...	71	HBC 2.9 $K^-p \rightarrow \Sigma/\Lambda K \bar{K}$

<sup>1</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations.

<sup>2</sup> Update of AKHMETSHIN 99D

<sup>3</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>4</sup> Using a total width of  $4.43 \pm 0.05$  MeV. Systematic uncertainty included.

<sup>5</sup> Using a total width of  $4.43 \pm 0.05$  MeV.

<sup>6</sup> PELLINEN 82 review includes AKERLOF 77, DAUM 81, BALDI 77, AYRES 74, DE-GROOT 74.

<sup>7</sup> Strongly correlated with AKHMETSHIN 04.

<sup>8</sup> Systematic errors not evaluated.

<sup>9</sup> Weighted and scaled average of 12 measurements of DIJKSTRA 86.

<sup>10</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

 $\phi(1020)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.266 ± 0.031 OUR AVERAGE</b>		Error includes scale factor of 1.2.		
4.29 ± 0.04 ± 0.07		<sup>1</sup> LEES	13Q	BABR $e^+e^- \rightarrow K^+K^-\gamma$
4.30 ± 0.06 ± 0.17	105k	AKHMETSHIN 06	CMD2	0.98–1.06 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
4.280 ± 0.033 ± 0.025	272k	<sup>2</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow K_L^0 K_S^0$
4.21 ± 0.04	1900k	<sup>3</sup> ACHASOV	01E	SND $e^+e^- \rightarrow K^+K^-, K_S K_L, \pi^+\pi^-\pi^0$
4.44 ± 0.09	55600	AKHMETSHIN 95	CMD2	$e^+e^- \rightarrow \text{hadrons}$
4.5 ± 0.7	1500	ARENTON	82	AEMS 11.8 polar. $pp \rightarrow K K$
4.2 ± 0.6	766	<sup>4</sup> IVANOV	81	OLYA 1–1.4 $e^+e^- \rightarrow K^+K^-$
4.3 ± 0.6		<sup>4</sup> CORDIER	80	DM1 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
4.36 ± 0.29	3681	<sup>4</sup> BUKIN	78C	OLYA $e^+e^- \rightarrow \text{hadrons}$
4.4 ± 0.6	984	<sup>4</sup> BESCH	74	CNTR 2 $\gamma p \rightarrow p K^+ K^-$
4.67 ± 0.72	681	<sup>4</sup> BALAKIN	71	OSPK $e^+e^- \rightarrow \text{hadrons}$
4.09 ± 0.29		BIZOT	70	OSPK $e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.37 ± 0.02		LEES	13F	BABR $D^+ \rightarrow K^+K^-\pi^+$
4.24 ± 0.02 ± 0.03	542k	<sup>5</sup> AKHMETSHIN 08	CMD2	1.02 $e^+e^- \rightarrow K^+K^-$
4.28 ± 0.13	12540	<sup>6</sup> AUBERT, B	05J	BABR $D^0 \rightarrow \bar{K}^0 K^+ K^-$
4.45 ± 0.06	271k	DIJKSTRA	86	SPEC 100 $\pi^- \text{Be}$
3.6 ± 0.8	337	<sup>4</sup> COOPER	78B	HBC 0.7–0.8 $\bar{p} p \rightarrow K_S^0 K_L^0 \pi^+\pi^-$
4.5 ± 0.50	1300	<sup>4,6</sup> AKERLOF	77	SPEC 400 $pA \rightarrow K^+K^-\pi^0$
4.5 ± 0.8	500	<sup>4,6</sup> AYRES	74	ASPK 3–6 $\pi^- p \rightarrow K^+K^-\pi^0, K^+K^-\pi^-, K^+K^-\Lambda/\Sigma^0$
3.81 ± 0.37		COSME	74B	OSPK $e^+e^- \rightarrow K_L^0 K_S^0$
3.8 ± 0.7	454	<sup>4</sup> BORENSTEIN	72	HBC 2.18 $K^-p \rightarrow K \bar{K} n$

<sup>1</sup> Using a phenomenological model based on KUHN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations.

<sup>2</sup> Update of AKHMETSHIN 99D

<sup>3</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S K_L$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.

<sup>4</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>5</sup> Strongly correlated with AKHMETSHIN 04.

<sup>6</sup> Systematic errors not evaluated.

 $\phi(1020)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $K^+K^-$	(48.9 ± 0.5 ) %	S=1.1
$\Gamma_2$ $K_L^0 K_S^0$	(34.2 ± 0.4 ) %	S=1.1
$\Gamma_3$ $\rho\pi^+ + \pi^+\pi^-\pi^0$	(15.32 ± 0.32 ) %	S=1.1
$\Gamma_4$ $\rho\pi^-$		
$\Gamma_5$ $\pi^+\pi^-\pi^0$		
$\Gamma_6$ $\eta\gamma$	( 1.309 ± 0.024 ) %	S=1.2
$\Gamma_7$ $\pi^0\gamma$	( 1.27 ± 0.06 ) × 10 <sup>-3</sup>	
$\Gamma_8$ $\ell^+\ell^-$	—	
$\Gamma_9$ $e^+e^-$	( 2.954 ± 0.030 ) × 10 <sup>-4</sup>	S=1.1
$\Gamma_{10}$ $\mu^+\mu^-$	( 2.87 ± 0.19 ) × 10 <sup>-4</sup>	
$\Gamma_{11}$ $\eta e^+ e^-$	( 1.08 ± 0.04 ) × 10 <sup>-4</sup>	
$\Gamma_{12}$ $\pi^+\pi^-$	( 7.4 ± 1.3 ) × 10 <sup>-5</sup>	
$\Gamma_{13}$ $\omega\pi^0$	( 4.7 ± 0.5 ) × 10 <sup>-5</sup>	
$\Gamma_{14}$ $\omega\gamma$	< 5 %	CL=84%
$\Gamma_{15}$ $\rho\gamma$	< 1.2 %	CL=90%
$\Gamma_{16}$ $\pi^+\pi^-\gamma$	( 4.1 ± 1.3 ) × 10 <sup>-5</sup>	
$\Gamma_{17}$ $f_0(980)\gamma$	( 3.22 ± 0.19 ) × 10 <sup>-4</sup>	S=1.1
$\Gamma_{18}$ $\pi^0\pi^0\gamma$	( 1.13 ± 0.06 ) × 10 <sup>-4</sup>	
$\Gamma_{19}$ $\pi^+\pi^-\pi^+\pi^-$	( 4.0 ± 2.8 ) × 10 <sup>-6</sup>	
$\Gamma_{20}$ $\pi^+\pi^+\pi^-\pi^-\pi^0$	< 4.6 %	CL=90%
$\Gamma_{21}$ $\pi^0 e^+ e^-$	( 1.12 ± 0.28 ) × 10 <sup>-5</sup>	
$\Gamma_{22}$ $\pi^0 \eta\gamma$	( 7.27 ± 0.30 ) × 10 <sup>-5</sup>	S=1.5
$\Gamma_{23}$ $a_0(980)\gamma$	( 7.6 ± 0.6 ) × 10 <sup>-5</sup>	
$\Gamma_{24}$ $K^0 \bar{K}^0 \gamma$	< 1.9 %	CL=90%
$\Gamma_{25}$ $\eta'(958)\gamma$	( 6.25 ± 0.21 ) × 10 <sup>-5</sup>	
$\Gamma_{26}$ $\eta\pi^0\pi^0\gamma$	< 2 %	CL=90%
$\Gamma_{27}$ $\mu^+\mu^-\gamma$	( 1.4 ± 0.5 ) × 10 <sup>-5</sup>	

$\Gamma_{28}^{\phantom{28}} \rho\gamma\gamma$	$< 1.2$	$\times 10^{-4}$	CL=90%
$\Gamma_{29}^{\phantom{29}} \eta\pi^+\pi^-$	$< 1.8$	$\times 10^{-5}$	CL=90%
$\Gamma_{30}^{\phantom{30}} \eta\mu^+\mu^-$	$< 9.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{31}^{\phantom{31}} \eta U \rightarrow \eta e^+e^-$	$< 1$	$\times 10^{-6}$	CL=90%

$$\Gamma_{32} \quad e^\pm \mu^\mp \quad LF \quad < 2 \quad \times 10^{-6} \quad \text{CL}=90\%$$

An overall fit to 30 branching ratios uses 79 measurements and one constraint to determine 14 parameters. The overall fit has a  $\chi^2 = 57.4$  for 66 degrees of freedom.

$x_2$	-72									
$x_3$	-53	-21								
$x_6$	-13	7	2							
$x_7$	-5	3	1	5						
$x_9$	30	-25	-10	-32	-15					
$x_{10}$	-4	3	1	3	2	-11				
$x_{12}$	-2	1	0	2	1	-5	1			
$x_{13}$	-2	2	1	2	1	-7	1	0		
$x_{17}$	0	0	0	0	0	0	0	0	0	
$x_{18}$	-6	4	2	17	3	-17	2	1	1	0
$x_{19}$	0	0	0	0	0	-1	0	0	0	0
$x_{23}$	0	0	0	0	0	0	0	0	0	0
$x_{25}$	-4	2	1	32	2	-10	1	1	1	0
	$x_1$	$x_2$	$x_3$	$x_6$	$x_7$	$x_9$	$x_{10}$	$x_{12}$	$x_{13}$	$x_{17}$

$x_{19}$	0		
$x_{23}$	0	0	
$x_{25}$	5	0	0
	$x_{18}$	$x_{19}$	$x_{23}$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
$58.9 \pm 0.5 \pm 2.4$	ACHASOV	00	SND $e^+ e^- \rightarrow \gamma \gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
$5.40 \pm 0.16^{+0.43}_{-0.40}$	ACHASOV	00	SND $e^+e^- \rightarrow \pi^0\gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
1.320±0.017±0.015	<sup>1</sup> AMBROSINO 05	KLOE	1.02 e <sup>+</sup> e <sup>-</sup> → μ <sup>+</sup> μ <sup>-</sup>

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>1.27 ± 0.04 OUR EVALUATION</b>			
<b>1.251 ± 0.021 OUR AVERAGE</b>	Error includes scale factor of 1.1.		
1.235 ± 0.006 ± 0.022	2 AKHMETSHIN 11	CMD2	1.02 e <sup>+</sup> e <sup>-</sup> → φ
1.32 ± 0.05 ± 0.03	3 AMBROSINO 00	KLOE	1.02 e <sup>+</sup> e <sup>-</sup> → φ
1.28 ± 0.05	AKHMETSHIN 95	CMD2	1.02 e <sup>+</sup> e <sup>-</sup> → φ

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>1.320±0.018±0.017</b>	AMBROSINO 05	KLOE	1.02 $e^+e^- \rightarrow \mu^+\mu^-$

- <sup>1</sup> Weighted average of  $\Gamma_{ee}$  and  $\sqrt{\Gamma_{ee}\Gamma_{\mu\mu}}$  from AMBROSINO 05 assuming lepton universality.
- <sup>2</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is  $0.99741 \pm 0.00007$ .
- <sup>3</sup> From forward-backward asymmetry and using  $\Gamma_{\text{total}} = 4.26 \pm 0.05$  MeV from the 2004 edition of this Review.

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.6340±0.0070±0.0039</b>	<sup>1</sup> LEES	13Q BABR	$e^+e^- \rightarrow K^+K^-\gamma$

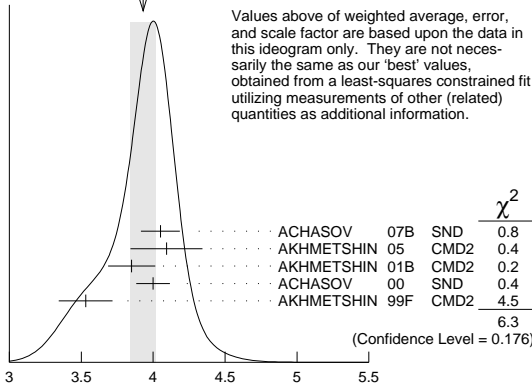
<sup>1</sup> Using a phenomenological model based on KUNN 90 with a sum of Breit-Wigner resonances for  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  and their higher mass excitations. The first error combines statistical and systematic uncertainties. The second one is due to the parametrization of the charged kaon form factor and mass calibration.

<u>VALUE (units 10<sup>-5</sup>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>14.46 ± 0.23 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>14.24 ± 0.30 OUR AVERAGE</b>				
14.27 ± 0.05 ± 0.31	542k	A KH METSHIN 08	CMD2	1.02 e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup>
13.93 ± 0.14 ± 0.99	1000k	<sup>1</sup> A CHASOV	01E SND	e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup> + K <sub>S</sub> K <sub>L</sub> π <sup>+</sup> π <sup>-</sup> π <sup>0</sup>

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>10.10<math>\pm</math>0.13 OUR FIT</b>				
<b>10.06<math>\pm</math>0.16 OUR AVERAGE</b>				
10.01 $\pm$ 0.04 $\pm$ 0.17	272k	<sup>2</sup> AKHMETSIN 04	CMD2	$e^+e^- \rightarrow K_S^0 K_S^0$
10.27 $\pm$ 0.07 $\pm$ 0.34	500k	<sup>1</sup> ACHASOV	01E SND	$e^+e^- \rightarrow K^+ K^-$ $K_S^0 K_S^0, \pi^+ \pi^- \pi^0$

$VALUE$ (units $10^{-5}$ )	$EVTS$	$DOCUMENT\ ID$	$TECN$	$COMMENT$
<b>4.53 <math>\pm</math> 0.10 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>4.46 <math>\pm</math> 0.12 OUR AVERAGE</b>				
4.51 $\pm$ 0.16 $\pm$ 0.11	105k	AKHMETSHIN 06	CMD2	$0.98-1.06\ e^+\pi^- \rightarrow \pi^+\pi^-\pi^0$
4.30 $\pm$ 0.08 $\pm$ 0.21		AUBERT,B	04N BABR	$10.6\ e^+\pi^- \rightarrow \pi^+\pi^-\pi^0\gamma$
4.665 $\pm$ 0.042 $\pm$ 0.261	400k	<sup>1</sup> ACHASOV	01E SND	$e^+\pi^- \rightarrow K^+K^-, K_S^0K_L, \pi^+\pi^-\pi^0$
4.35 $\pm$ 0.27 $\pm$ 0.08	11169	<sup>3</sup> AKHMETSHIN 98	CMD2	$e^+\pi^- \rightarrow \pi^+\pi^-\pi^0$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
4.38 $\pm$ 0.12		BENAYOUN 10	RVUE	$0.4-1.05\ e^+\pi^-$

<i>VALUE (units <math>10^{-6}</math>)</i>	<i>EVTs</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
<b>3.87 ± 0.07 OUR FIT</b>	Error	includes scale factor of 1.2.		
<b>3.93 ± 0.09 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
4.050 ± 0.067 ± 0.118	33k	<sup>4</sup> ACHASOV	07B	SND 0.6-1.38 $e^+e^- \rightarrow \eta\gamma$
4.093 ± 0.040 ± 0.043 ± 0.247	17.4k	<sup>5</sup> AKHMETSHIN	05	CMD2 0.60-1.38 $e^+e^- \rightarrow \eta\gamma$
3.850 ± 0.041 ± 0.159	23k	<sup>6,7</sup> AKHMETSHIN	01B	CMD2 $e^+e^- \rightarrow \eta\gamma$
4.00 ± 0.04 ± 0.11		<sup>8</sup> ACHASOV	00	SND $e^+e^- \rightarrow \eta\gamma$
3.53 ± 0.08 ± 0.17	2200	<sup>9,10</sup> AKHMETSHIN	99F	CMD2 $e^+e^- \rightarrow \eta\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
4.19 ± 0.06		<sup>11</sup> BENAYOUN	10	RVUE 0.4-1.05 $e^+e^-$



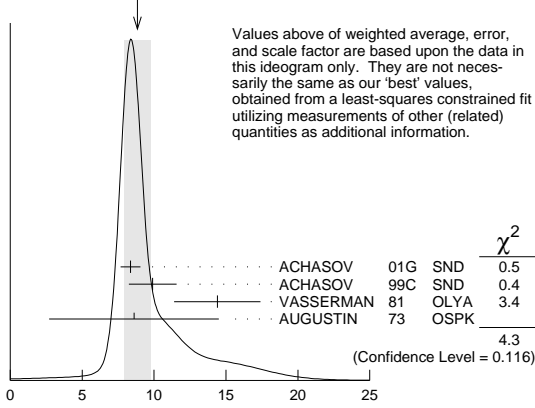
$$\Gamma(\eta\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_6/\Gamma \times \Gamma_9/\Gamma$$

See key on page 601

## Meson Particle Listings

 $\phi(1020)$  $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_7/\Gamma \times \Gamma_9/\Gamma$ VALUE (units  $10^{-7}$ ) EVTS DOCUMENT ID TECN COMMENT**3.74±0.18 OUR FIT****3.71±0.21 OUR AVERAGE**3.75±0.11±0.29 18680 AKHMETSHIN 05 CMD2 0.60-1.38  $e^+e^- \rightarrow \pi^0\gamma$   
3.67±0.10 $^{+0.27}_{-0.25}$  12 ACHASOV 00 SND  $e^+e^- \rightarrow \pi^0\gamma$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.29±0.11 11 BENAYOUN 10 RVUE 0.4-1.05  $e^+e^-$  $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_{10}/\Gamma \times \Gamma_9/\Gamma$ VALUE (units  $10^{-8}$ ) DOCUMENT ID TECN COMMENT**8.5 $^{+0.5}_{-0.6}$  OUR FIT****8.8±0.9 OUR AVERAGE** Error includes scale factor of 1.5. See the ideogram below.8.36±0.59±0.37 ACHASOV 01G SND  $e^+e^- \rightarrow \mu^+\mu^-$   
9.9±1.4±0.9 9 ACHASOV 99C SND  $e^+e^- \rightarrow \mu^+\mu^-$   
14.4±3.0 3 VASSERMAN 81 OLYA  $e^+e^- \rightarrow \mu^+\mu^-$   
8.6±5.9 3 AUGUSTIN 73 OSPK  $e^+e^- \rightarrow \mu^+\mu^-$ WEIGHTED AVERAGE  
8.8±0.9 (Error scaled by 1.5) $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_{10}/\Gamma \times \Gamma_9/\Gamma$  $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_{12}/\Gamma \times \Gamma_9/\Gamma$ VALUE (units  $10^{-8}$ ) DOCUMENT ID TECN COMMENT**2.2±0.4 OUR FIT****2.2±0.4 OUR AVERAGE**2.1±0.3±0.3 9 ACHASOV 00C SND  $e^+e^- \rightarrow \pi^+\pi^-$   
1.95 $^{+1.15}_{-0.87}$  3 GOLUBEV 86 ND  $e^+e^- \rightarrow \pi^+\pi^-$   
6.01 $^{+3.19}_{-2.51}$  3 VASSERMAN 81 OLYA  $e^+e^- \rightarrow \pi^+\pi^-$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.31±0.99 13 BENAYOUN 13 RVUE 0.4-1.05  $e^+e^-$  $\Gamma(\omega\pi^0)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_{13}/\Gamma \times \Gamma_9/\Gamma$ VALUE (units  $10^{-8}$ ) DOCUMENT ID TECN COMMENT**1.40±0.15 OUR FIT****1.37±0.17±0.01**14,15 AMBROSINO 08G KLOE  $e^+e^- \rightarrow \pi^+\pi^-\pi^0, 2\pi^0\gamma$  $\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_{18}/\Gamma \times \Gamma_9/\Gamma$ VALUE (units  $10^{-8}$ ) DOCUMENT ID TECN COMMENT**3.34±0.17 OUR FIT****3.33 $^{+0.04+0.19}_{-0.09-0.20}$**  16 AMBROSINO 07 KLOE  $e^+e^- \rightarrow \pi^0\pi^0\gamma$  $\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_{19}/\Gamma \times \Gamma_9/\Gamma$ VALUE (units  $10^{-9}$ ) EVTS DOCUMENT ID TECN COMMENT**1.2 $^{+0.8}_{-0.7}$  OUR FIT****1.17±0.52±0.64** 3285 9 AKHMETSHIN 00E CMD2  $e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$ 1 From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-$ ,  $K_S^0K_L^0$ ,  $\pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00B for the  $\eta\gamma$  decay mode.  
2 Update of AKHMETSHIN 99D  
3 Recalculated by us from the cross section in the peak.4 From a combined fit of  $\sigma(e^+e^- \rightarrow \eta\gamma)$  with  $\eta \rightarrow 3\pi^0$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$ , and fixing  $B(\eta \rightarrow 3\pi^0) / B(\eta \rightarrow \pi^+\pi^-\pi^0) = 1.44 \pm 0.04$ . Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.5 From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .  
6 From the  $\eta \rightarrow 3\pi^0$  decay and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .7 The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).  
8 From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow 2\gamma) = (39.21 \pm 0.34) \times 10^{-2}$ .

9 Recalculated by the authors from the cross section in the peak.

10 From the  $\eta \rightarrow \pi^+\pi^-\pi^0$  decay and using  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (23.1 \pm 0.5) \times 10^{-2}$ .11 A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-$ ,  $\pi^+\pi^-\pi^0$ ,  $\pi^0\gamma$ ,  $\eta\gamma$  data.12 From the  $\pi^0 \rightarrow 2\gamma$  decay and using  $B(\pi^0 \rightarrow 2\gamma) = (98.798 \pm 0.032) \times 10^{-2}$ .13 A simultaneous fit to  $e^+e^- \rightarrow \pi^+\pi^-$ ,  $\pi^+\pi^-\pi^0$ ,  $\pi^0\gamma$ ,  $\eta\gamma$ ,  $K\bar{K}$ , and  $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$  data.

14 Recalculated by the authors from the cross section at the peak.

15 AMBROSINO 08G reports  $[\Gamma(\phi(1020) \rightarrow \omega\pi^0)/\Gamma_{\text{total}} \times \Gamma(\phi(1020) \rightarrow e^+e^-)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow \pi^+\pi^-\pi^0)] = (1.22 \pm 0.13 \pm 0.08) \times 10^{-8}$  which we divide by our best value  $B(\omega(782) \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

16 Calculated by the authors from the cross section at the peak.

 $\phi(1020)$  BRANCHING RATIOS $\Gamma(K^+K^-)/\Gamma_{\text{total}} \quad \Gamma_1/\Gamma$ 

VALUE EVTS DOCUMENT ID TECN COMMENT

**0.489±0.005 OUR FIT** Error includes scale factor of 1.1.**0.493±0.010 OUR AVERAGE**0.492±0.012 2913 AKHMETSHIN 95 CMD2  $e^+e^- \rightarrow K^+K^-$   
0.44±0.05 321 KALBFLEISCH 76 HBC 2.18  $K^-p \rightarrow \Lambda K^+K^-$   
0.49±0.06 270 DEGROOT 74 HBC 4.2  $K^-p \rightarrow \Lambda\phi$   
0.540±0.034 565 BALAKIN 71 OSPK  $e^+e^- \rightarrow K^+K^-$   
0.48±0.04 252 LINDSEY 66 HBC 2.1-2.7  $K^-p \rightarrow \Lambda K^+K^-$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.493±0.003±0.007 1 AKHMETSHIN 11 CMD2 1.02  $e^+e^- \rightarrow K^+K^-$   
0.476±0.017 1000k 2 ACHASOV 01E SND  $e^+e^- \rightarrow K^+K^-$ ,  $K_S^0K_L^0$ ,  $\pi^+\pi^-\pi^0$  $\Gamma(K_L^0K_S^0)/\Gamma_{\text{total}} \quad \Gamma_2/\Gamma$ 

VALUE EVTS DOCUMENT ID TECN COMMENT

**0.342±0.004 OUR FIT** Error includes scale factor of 1.1.**0.331±0.009 OUR AVERAGE**0.335±0.010 40644 AKHMETSHIN 95 CMD2  $e^+e^- \rightarrow K_L^0K_S^0$   
0.326±0.035 DOLINSKY 91 ND  $e^+e^- \rightarrow K_L^0K_S^0$   
0.310±0.024 DRUZHININ 84 ND  $e^+e^- \rightarrow K_L^0K_S^0$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.336±0.002±0.006 1 AKHMETSHIN 11 CMD2 1.02  $e^+e^- \rightarrow K_S^0K_L^0$   
0.351±0.013 500k 2 ACHASOV 01E SND  $e^+e^- \rightarrow K^+K^-$ ,  $K_S^0K_L^0$ ,  $\pi^+\pi^-\pi^0$ 0.27±0.03 133 KALBFLEISCH 76 HBC 2.18  $K^-p \rightarrow \Lambda K_L^0K_S^0$ 0.257±0.030 95 BALAKIN 71 OSPK  $e^+e^- \rightarrow K_L^0K_S^0$ 0.40±0.04 167 LINDSEY 66 HBC 2.1-2.7  $K^-p \rightarrow \Lambda K_L^0K_S^0$  $\Gamma(K_L^0K_S^0)/\Gamma(K^+K^-) \quad \Gamma_2/\Gamma_1$ 

VALUE EVTS DOCUMENT ID TECN COMMENT

**0.698±0.014 OUR FIT** Error includes scale factor of 1.1.**0.740±0.031 OUR AVERAGE**0.70±0.06 2732 BUKIN 78C OLYA  $e^+e^- \rightarrow K_L^0K_S^0$   
0.82±0.08 LOSTY 78 HBC 4.2  $K^-p \rightarrow \phi$  hyperon  
0.71±0.05 LAVEN 77 HBC 10  $K^-p \rightarrow K^+K^-\Lambda$   
0.71±0.08 LYONS 77 HBC 3-4  $K^-p \rightarrow \Lambda\phi$   
0.89±0.10 144 AGUILAR-... 72B HBC 3.9,4.6  $K^-p$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.68±0.03 3 AKHMETSHIN 95 CMD2  $e^+e^- \rightarrow K_L^0K_S^0$ ,  $K^+K^-$  $\Gamma(K_L^0K_S^0)/\Gamma(K\bar{K}) \quad \Gamma_2/(\Gamma_1+\Gamma_2)$ 

VALUE EVTS DOCUMENT ID TECN COMMENT

**0.411±0.005 OUR FIT** Error includes scale factor of 1.1.**0.45±0.04 OUR AVERAGE**0.44±0.07 LONDON 66 HBC 2.24  $K^-p \rightarrow \Lambda K\bar{K}$   
0.48±0.07 52 BADIER 65B HBC 3  $K^-p$   
0.40±0.10 34 SCHLEIN 63 HBC 1.95  $K^-p \rightarrow \Lambda K\bar{K}$  $[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma_{\text{total}} \quad \Gamma_3/\Gamma$ 

VALUE EVTS DOCUMENT ID TECN COMMENT

**0.1532±0.0032 OUR FIT** Error includes scale factor of 1.1.**0.151±0.009 OUR AVERAGE** Error includes scale factor of 1.7.0.161±0.008 11761 AKHMETSHIN 95 CMD2  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$   
0.143±0.007 DOLINSKY 91 ND  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.155±0.002±0.005 1 AKHMETSHIN 11 CMD2 1.02  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$ 0.159±0.008 400k 2 ACHASOV 01E SND  $e^+e^- \rightarrow K^+K^-$ ,  $K_S^0K_L^0$ ,  $\pi^+\pi^-\pi^0$ 0.145±0.009±0.003 11169 4 AKHMETSHIN 98 CMD2  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$ 0.139±0.007 5 PARROUR 76B OSPK  $e^+e^-$  $[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma(K^+K^-) \quad \Gamma_3/\Gamma_1$ 

VALUE EVTS DOCUMENT ID TECN COMMENT

**0.313±0.009 OUR FIT** Error includes scale factor of 1.1.**0.28±0.09**34 AGUILAR-... 72B HBC 3.9,4.6  $K^-p$

# Meson Particle Listings

## $\phi(1020)$

### $[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma(K\bar{K})$ $\Gamma_3/(\Gamma_1+\Gamma_2)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.184±0.005 OUR FIT</b>	Error includes scale factor of 1.1.		
<b>0.24 ±0.04 OUR AVERAGE</b>			
0.237±0.039	CERRADA	77B HBC	4.2 $K^-p \rightarrow \Lambda 3\pi$
0.30 ±0.15	LONDON	66 HBC	2.24 $K^-p \rightarrow \Lambda \pi^+\pi^-\pi^0$

### $[\Gamma(\rho\pi) + \Gamma(\pi^+\pi^-\pi^0)]/\Gamma(K_S^0 K_S^0)$ $\Gamma_3/\Gamma_2$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.448±0.012 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>0.51 ±0.05 OUR AVERAGE</b>				
0.56 ±0.07	3681	BUKIN	78c OLYA	$e^+e^- \rightarrow K_L^0 K_S^0, \pi^+\pi^-\pi^0$
0.47 ±0.06	516	COSME	74 OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$

### $\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_5/\Gamma$

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
≈ 0.0087		1.98M	6,7 ALOISIO	03 KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<0.0006	90		8 ACHASOV	02 SND	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<0.23	90		8 CORDIER	80 DM1	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<0.20	90		8 PARROUR	76B OSPK	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$

### $\Gamma(\eta\gamma)/\Gamma_{\text{total}}$ $\Gamma_6/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.309±0.024 OUR FIT</b> Error includes scale factor of 1.2.				
<b>1.26 ±0.04 OUR AVERAGE</b>				
1.246±0.025±0.057	10k	<sup>9</sup> ACHASOV	98F SND	$e^+e^- \rightarrow 7\gamma$
1.18 ±0.11	279	<sup>10</sup> AKHMETSHIN	95 CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^0 3\gamma$
1.30 ±0.06		<sup>11</sup> DRUZHININ	84 ND	$e^+e^- \rightarrow 3\gamma$
1.4 ±0.2		<sup>12</sup> DRUZHININ	84 ND	$e^+e^- \rightarrow 6\gamma$
0.88 ±0.20	290	KURDADZE	83C OLYA	$e^+e^- \rightarrow 3\gamma$
1.35 ±0.29		ANDREWS	77 CNTR	6.7-10 $\gamma$ Cu
1.5 ±0.4	54	<sup>11</sup> COSME	76 OSPK	$e^+e^-$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1.38 ±0.02 ±0.02		<sup>1</sup> AKHMETSHIN	11 CMD2	1.02 $e^+e^- \rightarrow \eta\gamma$
1.37 ±0.05 ±0.01	33k	<sup>13</sup> ACHASOV	07B SND	0.6-1.38 $e^+e^- \rightarrow \eta\gamma$
1.373±0.014±0.085	17.4k	<sup>14,15</sup> AKHMETSHIN	05 CMD2	0.60-1.38 $e^+e^- \rightarrow \eta\gamma$
1.287±0.013±0.063		<sup>16,17</sup> AKHMETSHIN	01B CMD2	$e^+e^- \rightarrow \eta\gamma$
1.338±0.012±0.052		<sup>18</sup> ACHASOV	00 SND	$e^+e^- \rightarrow \eta\gamma$
1.18 ±0.03 ±0.06	2200	<sup>19</sup> AKHMETSHIN	99F CMD2	$e^+e^- \rightarrow \eta\gamma$
1.21 ±0.07		<sup>20</sup> BENAYOUN	96 RVUE	0.54-1.04 $e^+e^- \rightarrow \eta\gamma$

### $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$ $\Gamma_7/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.27 ±0.06 OUR FIT</b>				
<b>1.31 ±0.13 OUR AVERAGE</b>				
1.30 ±0.13		DRUZHININ	84 ND	$e^+e^- \rightarrow 3\gamma$
1.4 ±0.5	32	COSME	76 OSPK	$e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.258±0.037±0.077	18680	21,22 AKHMETSHIN	05 CMD2	0.60-1.38 $e^+e^- \rightarrow \pi^0\gamma$
1.226±0.036 $\pm^{+0.096}_{-0.089}$		23 ACHASOV	00 SND	$e^+e^- \rightarrow \pi^0\gamma$
1.26 ±0.17		20 BENAYOUN	96 RVUE	0.54-1.04 $e^+e^- \rightarrow \pi^0\gamma$

### $\Gamma(\eta\gamma)/\Gamma(\pi^0\gamma)$ $\Gamma_6/\Gamma_7$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
10.9±0.3 $\pm^{+0.7}_{-0.8}$	ACHASOV	00 SND	$e^+e^- \rightarrow \eta\gamma, \pi^0\gamma$

### $\Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.954±0.030 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>2.98 ±0.07 OUR AVERAGE</b>				
2.93 ±0.14	1900k	24 ACHASOV	01E SND	$e^+e^- \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0$
2.88 ±0.09	55600	AKHMETSHIN	95 CMD2	$e^+e^- \rightarrow$ hadrons
3.00 ±0.21	3681	BUKIN	78c OLYA	$e^+e^- \rightarrow$ hadrons
3.10 ±0.14		25 PARROUR	76 OSPK	$e^+e^-$
3.3 ±0.3		COSME	74 OSPK	$e^+e^- \rightarrow$ hadrons
2.81 ±0.25	681	BALAKIN	71 OSPK	$e^+e^- \rightarrow$ hadrons
3.50 ±0.27		CHATELUS	71 OSPK	$e^+e^-$

### $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.87±0.19 OUR FIT</b>			
<b>2.5 ±0.4 OUR AVERAGE</b>			
2.69±0.46	26 HAYES	71 CNTR	8.3,9.8 $\gamma$ C $\rightarrow \mu^+\mu^-X$
2.17±0.60	26 EARLES	70 CNTR	6.0 $\gamma$ C $\rightarrow \mu^+\mu^-X$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.87±0.20±0.14	27 ACHASOV	01G SND	$e^+e^- \rightarrow \mu^+\mu^-$
3.30±0.45±0.32	4 ACHASOV	99c SND	$e^+e^- \rightarrow \mu^+\mu^-$
4.83±1.02	28 VASSERMAN	81 OLYA	$e^+e^- \rightarrow \mu^+\mu^-$
2.87±1.98	28 AUGUSTIN	73 OSPK	$e^+e^- \rightarrow \mu^+\mu^-$

### $\Gamma(\eta e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.08 <math>\pm 0.04</math> OUR AVERAGE</b>				
1.075 $\pm 0.007 \pm 0.038$	30k	<sup>29</sup> BABUSCI	15 KLOE	1.02 $e^+e^- \rightarrow \eta e^+e^-$
1.19 $\pm 0.19 \pm 0.12$	213	<sup>30</sup> ACHASOV	01B SND	$e^+e^- \rightarrow \eta e^+e^-$
1.14 $\pm 0.10 \pm 0.06$	355	<sup>31</sup> AKH METSHIN 01	CMD2	$e^+e^- \rightarrow \eta e^+e^-$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1.13 $\pm 0.14 \pm 0.07$	183	<sup>32</sup> AKH METSHIN 01	CMD2	$e^+e^- \rightarrow \eta e^+e^-$
1.21 $\pm 0.14 \pm 0.09$	130	<sup>33</sup> AKH METSHIN 01	CMD2	$e^+e^- \rightarrow \eta e^+e^-$
1.04 $\pm 0.20 \pm 0.08$	42	<sup>34</sup> AKH METSHIN 01	CMD2	$e^+e^- \rightarrow \eta e^+e^-$
1.3 $\begin{smallmatrix} +0.8 \\ -0.6 \end{smallmatrix}$	7	GOLUBEV	85 ND	$e^+e^- \rightarrow \eta e^+e^-$

### $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.71±0.11±0.09		4 ACHASOV	00c ND	$e^+e^- \rightarrow \pi^+\pi^-$
0.65 $\pm^{+0.38}_{-0.29}$		4 GOLUBEV	86 ND	$e^+e^- \rightarrow \pi^+\pi^-$
2.01 $\pm^{+1.07}_{-0.84}$		4 VASSERMAN	81 OLYA	$e^+e^- \rightarrow \pi^+\pi^-$
<6.6	95	BUKIN	78B OLYA	$e^+e^- \rightarrow \pi^+\pi^-$
<2.7	95	ALVENSLEB...	72 CNTR	6.7 $\gamma$ C $\rightarrow$ $C\pi^+\pi^-$

### $\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{13}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.7±0.5 OUR FIT</b>			
<b>5.2<math>\pm^{+1.3}_{-1.1}</math></b>	35,36 AULCHENKO	00A SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.4±0.6	37 AMBROSINO	08G KLOE	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma, 2\pi^0\gamma$
~5.4	38 ACHASOV	00E SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
5.5 $\pm^{+1.6}_{-1.4}$ ±0.3	36,39 AULCHENKO	00A SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
4.8 $\pm^{+1.9}_{-1.7}$ ±0.8	38 ACHASOV	99 SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

### $\Gamma(\omega\gamma)/\Gamma_{\text{total}}$ $\Gamma_{14}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.05</b>	84	LINDSEY	66 HBC	2.1-2.7 $K^-p \rightarrow \Lambda \pi^+\pi^-$ neutrals

### $\Gamma(\rho\gamma)/\Gamma_{\text{total}}$ $\Gamma_{15}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.12</b>	90	40 AKHMETSHIN	99B CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 7	90	AKHMETSHIN	97c CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
<200	84	LINDSEY	66 HBC	2.1-2.7 $K^-p \rightarrow \Lambda \pi^+\pi^-$ neutrals

### $\Gamma(\pi^+\pi^-\gamma)/\Gamma_{\text{total}}$ $\Gamma_{16}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.41±0.12±0.04</b>	30175	41 AKHMETSHIN	99B CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.3	90	42 AKHMETSHIN	97c CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$	
<600	90	KALBFLEISCH	75 HBC	2.18 $K^-p \rightarrow \Lambda \pi^+\pi^-\gamma$	
< 70	90	COSME	74 OSPK	$e^+e^- \rightarrow \pi^+\pi^-\gamma$	
<400	90	LINDSEY	65 HBC	2.1-2.7 $K^-p \rightarrow \Lambda \pi^+\pi^-$ neutrals	

### $\Gamma(f_0(980)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{17}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.22±0.19 OUR FIT</b>					
<b>3.21±0.19 OUR AVERAGE</b>					
3.21 $\pm^{+0.03}_{-0.09}$ ±0.18		43 AMBROSINO	07 KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
2.90±0.21±1.54		44 AKHMETSHIN	99c CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma, \pi^0\pi^0\gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.47±0.21	2438	45 ALOISIO	02D KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
3.5 ±0.3 $\pm^{+1.3}_{-0.5}$	419	46,47 ACHASOV	00H SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
1.93±0.46±0.50	27188	48 AKHMETSHIN	99B CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$	
3.05±0.25±0.72	268	49 AKHMETSHIN	99c CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
1.5 ±0.5	268	50 AKHMETSHIN	99c CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	
3.42±0.30±0.36	164	46 ACHASOV	98I SND	$e^+e^- \rightarrow 5\gamma$	
< 1	90	51 AKHMETSHIN	97c CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$	
< 7	90	52 AKHMETSHIN	97c CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$	
< 20	90	DRUZHININ	87 ND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$	

### $\Gamma(f_0(980)\gamma)/\Gamma(\eta\gamma)$ $\Gamma_{17}/\Gamma_6$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.46±0.15 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>2.6 ±0.2 <math>\pm^{+0.8}_{-0.3}</math></b>	419	46 ACHASOV	00H SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$

See key on page 601

## Meson Particle Listings

 $\phi(1020)$ 

$\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.07 \pm 0.06</math> OUR AVERAGE</b>					
$1.07^{+0.01}_{-0.03} \pm 0.06$			<sup>53</sup> AMBROSINO	07 KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$1.08 \pm 0.17 \pm 0.09$		268	AKHMETSHIN 99c	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.09 \pm 0.03 \pm 0.05$		2438	ALOISIO	02d KLOE	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
$1.158 \pm 0.093 \pm 0.052$		419 47,54	ACHASOV	00H SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
<10		90	DRUZHININ	87 ND	$e^+e^- \rightarrow 5\gamma$

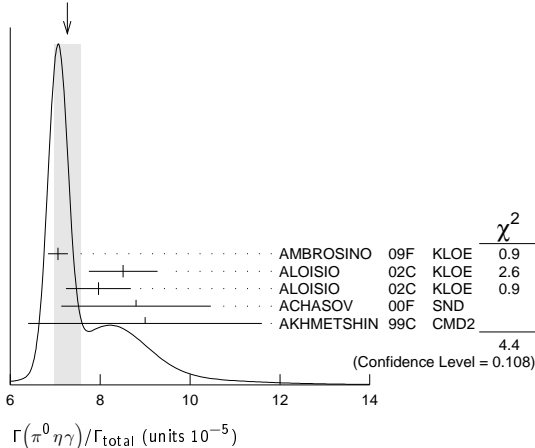
$\Gamma(\pi^0\pi^0\gamma)/\Gamma(\eta\gamma)$					$\Gamma_{18}/\Gamma_6$
VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.86 \pm 0.04</math> OUR FIT</b>					
<b><math>0.865 \pm 0.070 \pm 0.017</math></b>		419	<sup>54</sup> ACHASOV	00H SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.90 \pm 0.08 \pm 0.07$		164	ACHASOV	98i SND	$e^+e^- \rightarrow 5\gamma$

$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$3.93 \pm 1.74 \pm 2.14$		3285	AKHMETSHIN 00e	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
< 870		90	CORDIER	79 WIRE	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

$\Gamma(\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 4.6</b>		90	AKHMETSHIN 00e	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^-\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<150		95	BARKOV	88 CMD	$e^+e^- \rightarrow \pi^+\pi^-\pi^-\pi^0$

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.12 \pm 0.28</math> OUR AVERAGE</b>					
$1.01 \pm 0.28 \pm 0.29$		52	<sup>55</sup> ACHASOV	02d SND	$e^+e^- \rightarrow \pi^0 e^+ e^-$
$1.22 \pm 0.34 \pm 0.21$		46	<sup>56</sup> AKHMETSHIN 01c	CMD2	$e^+e^- \rightarrow \pi^0 e^+ e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<12		90	DOLINSKY	88 ND	$e^+e^- \rightarrow \pi^0 e^+ e^-$

$\Gamma(\pi^0\eta\gamma)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>7.27 \pm 0.30</math> OUR AVERAGE</b>					
$7.06 \pm 0.22$		16.9k	<sup>57</sup> AMBROSINO	09F KLOE	$1.02 e^+e^- \rightarrow \eta\pi^0\gamma$
$8.51 \pm 0.51 \pm 0.57$		607	<sup>58</sup> ALOISIO	02c KLOE	$e^+e^- \rightarrow \eta\pi^0\gamma$
$7.96 \pm 0.60 \pm 0.40$		197	<sup>59</sup> ALOISIO	02c KLOE	$e^+e^- \rightarrow \eta\pi^0\gamma$
$8.8 \pm 1.4 \pm 0.9$		36	<sup>60</sup> ACHASOV	00F SND	$e^+e^- \rightarrow \eta\pi^0\gamma$
$9.0 \pm 2.4 \pm 1.0$		80	AKHMETSHIN 99c	CMD2	$e^+e^- \rightarrow \eta\pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$7.01 \pm 0.10 \pm 0.20$		13.3k	<sup>58,61</sup> AMBROSINO	09F KLOE	$1.02 e^+e^- \rightarrow \eta\pi^0\gamma$
$7.12 \pm 0.13 \pm 0.22$		3.6k	<sup>59,62</sup> AMBROSINO	09F KLOE	$1.02 e^+e^- \rightarrow \eta\pi^0\gamma$
$8.3 \pm 2.3 \pm 1.2$		20	ACHASOV	98B SND	$e^+e^- \rightarrow 5\gamma$
<250		90	DOLINSKY	91 ND	$e^+e^- \rightarrow \pi^0\eta\gamma$

WEIGHTED AVERAGE  
 $7.27 \pm 0.30$  (Error scaled by 1.5)

$\Gamma(a_0(980)\gamma)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>7.6 \pm 0.6</math> OUR FIT</b>					
<b><math>7.6 \pm 0.6</math> OUR AVERAGE</b>					
$7.4 \pm 0.7$			<sup>63</sup> ALOISIO	02c KLOE	$e^+e^- \rightarrow \eta\pi^0\gamma$
$8.8 \pm 1.7$		36	<sup>64</sup> ACHASOV	00F SND	$e^+e^- \rightarrow \eta\pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$11 \pm 2$			<sup>65</sup> GOKALP	02 RVUE	$e^+e^- \rightarrow \eta\pi^0\gamma$
<500		90	DOLINSKY	91 ND	$e^+e^- \rightarrow \pi^0\eta\gamma$

$\Gamma(f_0(980)\gamma)/\Gamma(a_0(980)\gamma)$					$\Gamma_{17}/\Gamma_{23}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>6.1 \pm 0.6</math></b>		66	ALOISIO	02c KLOE	$e^+e^- \rightarrow \eta\pi^0\gamma$

$\Gamma(K^0\bar{K}^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; <math>1.9 \times 10^{-8}</math></b>		90	AMBROSINO	09c KLOE	$e^+e^- \rightarrow K_S^0\bar{K}_S^0\gamma$

$\Gamma(\eta'(958)\gamma)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.25 \pm 0.21</math> OUR FIT</b>					
<b><math>6.25 \pm 0.30</math> OUR AVERAGE</b>					
$6.25 \pm 0.28 \pm 0.11$		3407	<sup>67</sup> AMBROSINO	07A KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\gamma$

$6.7^{+2.8}_{-2.4} \pm 0.8$		12	<sup>68</sup> AULCHENKO	03B SND	$e^+e^- \rightarrow \eta'\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					

$6.7^{+5.0}_{-4.2} \pm 1.5$		7	AULCHENKO	03B SND	$e^+e^- \rightarrow 7\gamma$
$6.10 \pm 0.61 \pm 0.43$		120	<sup>69</sup> ALOISIO	02E KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\gamma$
$8.2^{+2.1}_{-1.9} \pm 1.1$		21	<sup>70</sup> AKHMETSHIN 00B	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
$4.9^{+2.2}_{-1.8} \pm 0.6$		9	<sup>71</sup> AKHMETSHIN 00F	CMD2	$e^+e^- \rightarrow \eta'(\pi^+\pi^-) \geq 2\gamma$
$6.4 \pm 1.6$		30	<sup>72</sup> AKHMETSHIN 00F	CMD2	$e^+e^- \rightarrow \eta'(\pi^+\pi^-) \gamma$
$6.7^{+3.4}_{-2.9} \pm 1.0$		5	<sup>73</sup> AULCHENKO	99 SND	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
<11		90	AULCHENKO	98 SND	$e^+e^- \rightarrow 7\gamma$
$12^{+7}_{-5} \pm 2$		6	<sup>70</sup> AKHMETSHIN 97B	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
<41		90	DRUZHININ	87 ND	$e^+e^- \rightarrow \gamma\eta\pi^+\pi^-$

$\Gamma(\eta'(958)\gamma)/\Gamma(K_L^0 K_S^0)$					$\Gamma_{25}/\Gamma_2$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.83 \pm 0.06</math> OUR FIT</b>					
<b><math>1.46^{+0.64}_{-0.54} \pm 0.18</math></b>		9	<sup>74</sup> AKHMETSHIN 00F	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^- \geq 2\gamma$

$\Gamma(\eta'(958)\gamma)/\Gamma(\eta\gamma)$					$\Gamma_{25}/\Gamma_6$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.77 \pm 0.15</math> OUR FIT</b>					
<b><math>4.78 \pm 0.20</math> OUR AVERAGE</b>					
$4.77 \pm 0.09 \pm 0.19$		3407	AMBROSINO	07A KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\gamma$
$4.70 \pm 0.47 \pm 0.31$		120	<sup>75</sup> ALOISIO	02E KLOE	$1.02 e^+e^- \rightarrow \pi^+\pi^-\gamma$
$6.5^{+1.7}_{-1.5} \pm 0.8$		21	AKHMETSHIN 00B	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					

$9.5^{+5.2}_{-4.0} \pm 1.4$		6	<sup>76</sup> AKHMETSHIN 97B	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
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$\Gamma(\eta\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 2</b>		90	AULCHENKO	98 SND	$e^+e^- \rightarrow 7\gamma$

$\Gamma(\mu^+\mu^-\gamma)/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.43 \pm 0.45 \pm 0.14</math></b>		27188	<sup>48</sup> AKHMETSHIN 99B	CMD2	$e^+e^- \rightarrow \mu^+\mu^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$2.3 \pm 1.0$		824 ± 33	<sup>77</sup> AKHMETSHIN 97C	CMD2	$e^+e^- \rightarrow \mu^+\mu^-\gamma$

$\Gamma(\rho\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 1.2</b>		90	AULCHENKO	08 CMD2	$\phi \rightarrow \pi^+\pi^-\gamma\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<5		90	AKHMETSHIN	98 CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma\gamma$

$\Gamma(\eta\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 1.8</b>		90	AKHMETSHIN 00E	CMD2	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 6.1		90	AULCHENKO	08 CMD2	$\phi \rightarrow \eta\pi^+\pi^-$
<30		90	AKHMETSHIN	98 CMD2	$e^+e^- \rightarrow \pi^+\pi^-\gamma\gamma$

# Meson Particle Listings

## $\phi(1020)$

$\Gamma(\eta\mu^+\mu^-)/\Gamma_{\text{total}}$		$\Gamma_{30}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;9.4</b>	90	AKHMETSHIN 01	CMD2	$e^+e^- \rightarrow \eta e^+e^-$	

$\Gamma(\eta U \rightarrow \eta e^+e^-)/\Gamma_{\text{total}}$		$\Gamma_{31}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1 <math>\times 10^{-6}</math></b>	90	78 BABUSCI	13B KLOE	1.02 $e^+e^- \rightarrow \eta e^+e^-$	

- <sup>1</sup> Combined analysis of the CMD-2 data on  $\phi \rightarrow K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0, \eta\gamma$  assuming that the sum of their branching fractions is 0.99741  $\pm$  0.00007.
- <sup>2</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.93 \pm 0.14) \times 10^{-4}$ .
- <sup>3</sup> Theoretical analysis of BRAMON 00 taking into account phase-space difference, electromagnetic radiative corrections, as well as isospin breaking, predicts 0.62. FLOREZ-BAEZ 08 predicts 0.63 considering also structure-dependent radiative corrections. FISCHBACH 02 calculates additional corrections caused by the close threshold and predicts 0.68. See also BENAYOUN 01 and DUBYNISKIY 07. BENAYOUN 12 obtains  $0.71 \pm 0.01$  in the HLS model.
- <sup>4</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .
- <sup>5</sup> Using  $\Gamma(\phi) = 4.1$  MeV. If interference between the  $\rho\pi$  and  $3\pi$  modes is neglected, the fraction of the  $\rho\pi$  is more than 80% at the 90% confidence level.
- <sup>6</sup> From a fit without limitations on charged and neutral  $\rho$  masses and widths.
- <sup>7</sup> Adding the direct and  $\omega\pi$  contributions and considering the interference between the  $\rho\pi$  and  $\pi^+\pi^-\pi^0$ .
- <sup>8</sup> Neglecting the interference between the  $\rho\pi$  and  $\pi^+\pi^-\pi^0$ .
- <sup>9</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$  and  $B(\eta \rightarrow 3\pi^0) = (32.2 \pm 0.4) \times 10^{-2}$ .
- <sup>10</sup> From  $\pi^+\pi^-\pi^0$  decay mode of  $\eta$ .
- <sup>11</sup> From  $2\gamma$  decay mode of  $\eta$ .
- <sup>12</sup> From  $3\pi^0$  decay mode of  $\eta$ .
- <sup>13</sup> ACHASOV 07b reports  $[\Gamma(\phi(1020) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\phi(1020) \rightarrow e^+e^-)] = (4.050 \pm 0.067 \pm 0.118) \times 10^{-6}$  which we divide by our best value  $B(\phi(1020) \rightarrow e^+e^-) = (2.954 \pm 0.030) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00b and ACHASOV 06a.
- <sup>14</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.98 \pm 0.04) \times 10^{-4}$  and  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .
- <sup>15</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .
- <sup>16</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$  and  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .
- <sup>17</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770), \omega(782), \phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).
- <sup>18</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .
- <sup>19</sup> From  $\pi^+\pi^-\pi^0$  decay mode of  $\eta$  and using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .
- <sup>20</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.
- <sup>21</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.98 \pm 0.04) \times 10^{-4}$ .
- <sup>22</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .
- <sup>23</sup> From the  $\pi^0 \rightarrow 2\gamma$  decay and using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .
- <sup>24</sup> From the combined fit assuming that the total  $\phi(1020)$  production cross section is saturated by those of  $K^+K^-, K_S^0 K_L^0, \pi^+\pi^-\pi^0$ , and  $\eta\gamma$  decays modes and using ACHASOV 00b for the  $\eta\gamma$  decay mode.
- <sup>25</sup> Using total width 4.2 MeV. They detect  $3\pi$  mode and observe significant interference with  $\omega$  tail. This is accounted for in the result quoted above.
- <sup>26</sup> Neglecting interference between resonance and continuum.
- <sup>27</sup> Using  $B(\phi \rightarrow e^+e^-) = (2.91 \pm 0.07) \times 10^{-4}$ .
- <sup>28</sup> Recalculated by us using  $B(\phi \rightarrow e^+e^-) = (2.99 \pm 0.08) \times 10^{-4}$ .
- <sup>29</sup> Using  $B(\eta \rightarrow 3\pi^0) = (32.57 \pm 0.23)\%$  from PDG 12.
- <sup>30</sup> Using  $B(\eta \rightarrow \gamma\gamma) = (39.25 \pm 0.32)\%$ ,  $B(\phi \rightarrow \eta\gamma) = (1.26 \pm 0.06)\%$ , and  $B(\phi \rightarrow e^+e^-) = (3.00 \pm 0.06) \times 10^{-4}$ .
- <sup>31</sup> The average of the branching ratios separately obtained from the  $\eta \rightarrow \gamma\gamma, 3\pi^0, \pi^+\pi^-\pi^0$  decays.
- <sup>32</sup> From  $\eta \rightarrow \gamma\gamma$  decays and using  $B(\eta \rightarrow \gamma\gamma) = (39.33 \pm 0.25) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 11) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .
- <sup>33</sup> From  $\eta \rightarrow 3\pi^0$  decays and using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.798 \pm 0.033) \times 10^{-2}$ ,  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 0.11) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .
- <sup>34</sup> From  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays and using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.798 \pm 0.033) \times 10^{-2}$ ,  $B(\pi^0 \rightarrow e^+e^-\gamma) = (1.198 \pm 0.032) \times 10^{-2}$ ,  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (23.0 \pm 0.4) \times 10^{-2}$ ,  $B(\phi \rightarrow \pi^+\pi^-\pi^0) = (15.5 \pm 0.6) \times 10^{-2}$ , and  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ .
- <sup>35</sup> Using the 1996 and 1998 data.
- <sup>36</sup>  $(2.3 \pm 0.3)\%$  correction for other decay modes of the  $\omega(782)$  applied.
- <sup>37</sup> Not independent of the corresponding  $\Gamma(\omega\pi^0) \times \Gamma(e^+e^-) / \Gamma_{\text{total}}^2$ .
- <sup>38</sup> Using the 1996 data.
- <sup>39</sup> Using the 1998 data.
- <sup>40</sup> Supersedes AKHMETSHIN 97c.
- <sup>41</sup> For  $E_\gamma > 20$  MeV and assuming that  $B(\phi(1020) \rightarrow f_0(980)\gamma)$  is negligible. Supersedes AKHMETSHIN 97c.
- <sup>42</sup> For  $E_\gamma > 20$  MeV and assuming that  $B(\phi(1020) \rightarrow f_0(980)\gamma)$  is negligible.
- <sup>43</sup> Obtained by the authors taking into account the  $\pi^+\pi^-\pi^0$  decay mode. Includes a component due to  $\pi\pi$  production via the  $f_0(500)$  meson. Supersedes ALOISIO 02b.
- <sup>44</sup> From the combined fit of the photon spectra in the reactions  $e^+e^- \rightarrow \pi^+\pi^-\gamma, \pi^0\pi^0\gamma$ .
- <sup>45</sup> From the negative interference with the  $f_0(500)$  meson of AITALA 01b using the ACHASOV 89 parameterization for the  $f_0(980)$ , a Breit-Wigner for the  $f_0(500)$ , and ACHASOV 01f for the  $\rho\pi$  contribution. Superseded by AMBROSINO 07.
- <sup>46</sup> Assuming that the  $\pi^0\pi^0\gamma$  final state is completely determined by the  $f_0\gamma$  mechanism, neglecting the decay  $B(\phi \rightarrow K\bar{K}\gamma)$  and using  $B(f_0 \rightarrow \pi^+\pi^-) = 2B(f_0 \rightarrow \pi^0\pi^0)$ .
- <sup>47</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.338 \pm 0.053) \times 10^{-2}$ .
- <sup>48</sup> For  $E_\gamma > 20$  MeV. Supersedes AKHMETSHIN 97c.
- <sup>49</sup> Neglecting other intermediate mechanisms ( $\rho\pi, \sigma\gamma$ ).

- <sup>50</sup> A narrow pole fit taking into account  $f_0(980)$  and  $f_0(1200)$  intermediate mechanisms.
- <sup>51</sup> For destructive interference with the Bremsstrahlung process
- <sup>52</sup> For constructive interference with the Bremsstrahlung process
- <sup>53</sup> Supersedes ALOISIO 02b.
- <sup>54</sup> Supersedes ACHASOV 98i. Excluding  $\omega\pi^0$ .
- <sup>55</sup> Using various branching ratios from the 2000 Edition of this Review (PDG 00).
- <sup>56</sup> Using  $B(\pi^0 \rightarrow \gamma\gamma) = 0.98798 \pm 0.00032$ ,  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033) \times 10^{-2}$ , and  $B(\eta \rightarrow \pi^+\pi^-\gamma) = (4.75 \pm 0.11) \times 10^{-2}$ .
- <sup>57</sup> Combined results of  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$  decay modes measurements.
- <sup>58</sup> From the decay mode  $\eta \rightarrow \gamma\gamma$ .
- <sup>59</sup> From the decay mode  $\eta \rightarrow \pi^+\pi^-\pi^0$ .
- <sup>60</sup> Supersedes ACHASOV 98b.
- <sup>61</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.304 \pm 0.025)\%$ ,  $B(\eta \rightarrow 3\pi^0) = (32.56 \pm 0.23)\%$ , and  $B(\eta \rightarrow \gamma\gamma) = (39.31 \pm 0.20)\%$ .
- <sup>62</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.304 \pm 0.025)\%$ ,  $B(\eta \rightarrow 3\pi^0) = (32.56 \pm 0.23)\%$ , and  $B(\eta \rightarrow \pi^+\pi^-\pi^0) = (22.73 \pm 0.28)\%$ .
- <sup>63</sup> Using  $M_{a_0(980)} = 984.8$  MeV and assuming  $a_0(980)\gamma$  dominance.
- <sup>64</sup> Assuming  $a_0(980)\gamma$  dominance in the  $\eta\pi^0\gamma$  final state.
- <sup>65</sup> Using data of ACHASOV 00f.
- <sup>66</sup> Using results of ALOISIO 02b and assuming that  $f_0(980)$  decays into  $\pi\pi$  only and  $a_0(980)$  into  $\eta\pi$  only.
- <sup>67</sup> AMBROSINO 07a reports  $[\Gamma(\phi(1020) \rightarrow \eta'(958)\gamma)/\Gamma_{\text{total}}] / [B(\phi(1020) \rightarrow \eta\gamma)] = (4.77 \pm 0.09 \pm 0.19) \times 10^{-3}$  which we multiply by our best value  $B(\phi(1020) \rightarrow \eta\gamma) = (1.309 \pm 0.024) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>68</sup> Averaging AULCHENKO 03b with AULCHENKO 99.
- <sup>69</sup> Using  $B(\phi \rightarrow \eta\gamma) = (1.297 \pm 0.033)\%$ .
- <sup>70</sup> Using the value  $B(\phi \rightarrow \eta\gamma) = (1.26 \pm 0.06) \times 10^{-2}$ .
- <sup>71</sup> Using  $B(\phi \rightarrow K_L^0 K_S^0) = (33.8 \pm 0.6)\%$ .
- <sup>72</sup> Averaging AKHMETSHIN 00b with AKHMETSHIN 00f.
- <sup>73</sup> Using the value  $B(\eta' \rightarrow \eta\pi^+\pi^-) = (43.7 \pm 1.5) \times 10^{-2}$  and  $B(\eta \rightarrow \gamma\gamma) = (39.25 \pm 0.31) \times 10^{-2}$ .
- <sup>74</sup> Using various branching ratios of  $K_S^0, K_L^0, \eta, \eta'$  from the 2000 edition (The European Physical Journal **C15** 1 (2000)) of this Review.
- <sup>75</sup> From the decay mode  $\eta' \rightarrow \eta\pi^+\pi^-, \eta \rightarrow \gamma\gamma$ .
- <sup>76</sup> Superseded by AKHMETSHIN 00b.
- <sup>77</sup> For  $E_\gamma > 20$  MeV.
- <sup>78</sup> For a narrow vector  $U$  with mass between 5 and 470 MeV, from the combined analysis of  $\eta \rightarrow \pi^+\pi^-\pi^0$  and  $\eta \rightarrow \pi^0\pi^0\pi^0$  from ARCHILLI 12. Measured 90% CL limits as a function of  $m_U$  range from  $2.2 \times 10^{-8}$  to  $10^{-6}$ .

### Lepton Family number (LF) violating modes

$\Gamma(e^\pm\mu^\mp)/\Gamma_{\text{total}}$		$\Gamma_{32}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2 <math>\times 10^{-6}</math></b>	90	ACHASOV	10A SND	$e^+e^- \rightarrow e^\pm\mu^\mp$	

### $\pi^+\pi^-\pi^0 / \rho\pi$ AMPLITUDE RATIO $a_1$ IN DECAY OF $\phi \rightarrow \pi^+\pi^-\pi^0$

NIECKNIG 12 describes final-state interactions between the three pions in a dispersive framework using data on the  $\pi\pi$   $P$ -wave scattering phase shift.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.1<math>\pm</math>1.2 OUR AVERAGE</b>					
10.1 $\pm$ 4.4 $\pm$ 1.7	80k		<sup>1</sup> AKHMETSHIN 06	CMD2	1.017-1.021 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
9.0 $\pm$ 1.1 $\pm$ 0.6	1.98M		<sup>2,3</sup> ALOISIO	03 KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- $-6 < a_1 < 6$  500k <sup>3</sup> ACHASOV 02 SND  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
- $-16 < a_1 < 11$  90 9.8k <sup>1,4</sup> AKHMETSHIN 98 CMD2  $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
- <sup>1</sup> Dalitz plot analysis taking into account interference between the contact and  $\rho\pi$  amplitudes.
- <sup>2</sup> From a fit without limitations on charged and neutral  $\rho$  masses and widths.
- <sup>3</sup> Recalculated by us to match the notations of AKHMETSHIN 98.
- <sup>4</sup> Assuming zero phase for the contact term.

### PARAMETER $\beta$ IN $\phi \rightarrow \eta e^+e^-$ DECAY

In the one-pole approximation the electromagnetic transition form factor for  $\phi \rightarrow \eta e^+e^-$  is given as a function of the  $e^+e^-$  invariant mass squared,  $q^2$ , by the expression:

$$|F(q^2)|^2 = (1 - q^2/\Lambda^2)^{-2},$$

where vector meson dominance predicts parameter  $\Lambda \approx 0.770$  GeV ( $\Lambda^{-2} \approx 1.687$  GeV $^{-2}$ ). The slope of this form factor,  $\beta = dF/dq^2(q^2=0)$ , equals  $\Lambda^{-2}$  in this approximation.

The measurements below obtain  $\beta$  in the one-pole approximation.

VALUE (GeV $^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.29<math>\pm</math>0.13 OUR AVERAGE</b>					
1.28 $\pm$ 0.10 $^{+0.09}_{-0.08}$	30k		BABUSCI	15 KLOE	1.02 $e^+e^- \rightarrow \eta e^+e^-$
3.8 $\pm$ 1.8	213		<sup>1</sup> ACHASOV	01B SND	1.02 $e^+e^- \rightarrow \eta e^+e^-$

- <sup>1</sup> The uncertainty is statistical only. The systematic one is negligible, in comparison.



See key on page 601

## Meson Particle Listings

 $\phi(1020)$ ,  $h_1(1170)$  $\phi(1020)$  REFERENCES

BABUSCI	15	PL B742 1	D. Babusci <i>et al.</i>	(KLOE-2 Collab.)
BABUSCI	13B	PL B720 111	D. Babusci <i>et al.</i>	(KLOE-2 Collab.)
BENAYOUN	13	EPJ C73 2453	M. Benayoun, P. David, L. DeBuono	(PARIN, BERLIN+)
LEES	13F	PR D87 052010	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13Q	PR D88 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ARCHILLI	12	PL B706 251	F. Archilli <i>et al.</i>	(KLOE-2 Collab.)
BENAYOUN	12	EPJ C72 1848	M. Benayoun <i>et al.</i>	
NIECKNIG	12	EPJ C72 2014	F. Niecknig, B. Kubis, S.P. Schneider	(BONN)
PDG	12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
AKHMETSHIN	11	PL B695 412	R. Akhmetshin <i>et al.</i>	(CMD2 Collab.)
ACHASOV	10A	PR D81 057102	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
BENAYOUN	10	EPJ C65 211	M. Benayoun <i>et al.</i>	
AMBROSINO	09C	PL B679 10	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AMBROSINO	09F	PL B681 5	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AKHMETSHIN	08	PL B669 217	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AMBROSINO	08G	PL B669 223	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AULCHENKO	08	JETPL 88 85	V. Aulchenko <i>et al.</i>	(CMD-2 Collab.)
Translated from ZETFP 88 93.				
FLOREZ-BAEZ	08	PR D78 077301	F.V. Florez-Baez, G. Lopez Castro	
ACHASOV	07B	PR D76 077101	M.N. Achasov <i>et al.</i>	(SND Collab.)
AMBROSINO	07	EPJ C49 473	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AMBROSINO	07A	PL B648 267	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
DUBYNSKIY	07	PR D75 113001	S. Dubynskiy <i>et al.</i>	
ACHASOV	06A	PR D74 014016	M.N. Achasov <i>et al.</i>	(SND Collab.)
AKHMETSHIN	06	PL B642 203	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AKHMETSHIN	05	PL B605 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AMBROSINO	05	PL B608 199	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AUBERT,B	05J	PR D72 052008	B. Aubert <i>et al.</i>	(BABAR Collab.)
AKHMETSHIN	04	PL B578 285	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AUBERT,B	04N	PR D70 072004	B. Aubert <i>et al.</i>	(BABAR Collab.)
ALOISIO	03	PL B561 55	A. Aloisio <i>et al.</i>	(KLOE Collab.)
AULCHENKO	03B	JETP 97 24	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
Translated from ZETFP 124 28.				
ACHASOV	02	PR D65 032002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	02D	JETPL 75 449	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
Translated from ZETFP 75 539.				
ALOISIO	02C	PL B536 209	A. Aloisio <i>et al.</i>	(KLOE Collab.)
ALOISIO	02	PL B537 21	A. Aloisio <i>et al.</i>	(KLOE Collab.)
ALOISIO	02E	PL B541 45	A. Aloisio <i>et al.</i>	(KLOE Collab.)
FISCHBACH	02	PL B526 355	E. Fischbach, A.W. Overhauser, B. Woodahl	
GOKALP	02	JP G28 2703	A. Gokalp <i>et al.</i>	
ACHASOV	01B	PL B504 275	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	01E	PR D63 072002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	01F	PR D63 094007	M.N. Achasov, V.V. Gubin	(Novosibirsk SND Collab.)
ACHASOV	01G	PRL 86 1698	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AITALA	01B	PRL 86 770	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AKHMETSHIN	01	PL B501 191	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	01B	PL B509 217	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	01C	PL B503 237	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BENAYOUN	01	EPJ C22 503	M. Benayoun, H.B. O'Connell	
ACHASOV	00	EPJ C12 25	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00B	JETP 90 17	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
Translated from ZETFP 117 22.				
ACHASOV	00C	PL B474 188	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00D	JETPL 72 282	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
Translated from ZETFP 72 411.				
ACHASOV	00E	NP B569 158	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00F	PL B479 53	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00H	PL B485 349	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	00B	PL B473 337	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	00E	PL B491 81	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	00F	PL B494 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	00A	JETP 90 927	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
Translated from ZETFP 117 1067.				
BRAMON	00	PL B486 406	A. Bramon <i>et al.</i>	
PDG	00	EPJ C15 1	D.E. Groom <i>et al.</i>	(PDG Collab.)
ACHASOV	99	PL B449 122	M.N. Achasov <i>et al.</i>	
ACHASOV	99C	PL B456 304	M.N. Achasov <i>et al.</i>	
AKHMETSHIN	99B	PL B462 371	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99C	PL B462 380	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99D	PL B466 385	R.R. Akhmetshin <i>et al.</i>	
Also				
AKHMETSHIN	99F	PL B508 217 (errat.)	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	99C	PL B460 242	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AULCHENKO	99	JETPL 69 97	V.M. Aulchenko <i>et al.</i>	
Translated from ZETFP 69 87.				
ACHASOV	98B	PL B438 441	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98F	JETPL 68 573	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98H	PL B440 442	M.N. Achasov <i>et al.</i>	
AKHMETSHIN	98	PL B434 426	R.R. Akhmetshin <i>et al.</i>	(CMD-2 Collab.)
AULCHENKO	98	PL B436 199	V.M. Aulchenko <i>et al.</i>	(Novosibirsk SND Collab.)
BARBERIS	98	PL B432 436	D. Barberis <i>et al.</i>	(Omega E xpt.)
AKHMETSHIN	97B	PL B415 445	R.R. Akhmetshin <i>et al.</i>	(NOVO, BOST, PITT+)
AKHMETSHIN	97C	PL B415 452	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BENAYOUN	96	ZPHY C72 221	M. Benayoun <i>et al.</i>	(IPNP, NOVO)
AKHMETSHIN	95	PL B364 199	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ACHASOV	89	NP B315 465	M.N. Achasov, V.N. Ivanchenko	
DOLINSKY	89	ZPHY C42 511	S.I. Dolinsky <i>et al.</i>	(NOVO)
BARKOV	88	SJNP 47 248	L.M. Barkov <i>et al.</i>	(NOVO)
Translated from YAF 47 333.				
DOLINSKY	88	SJNP 48 277	S.I. Dolinsky <i>et al.</i>	(NOVO)
Translated from YAF 48 442.				
DRUZHININ	87	ZPHY C37 1	V.P. Druzhinin <i>et al.</i>	(NOVO)
ARMSTRONG	86	PL B668 245	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
ATKINSON	86	ZPHY C30 521	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
BEBEK	86	PRL 56 1893	C. Bebek <i>et al.</i>	(CLEO Collab.)
DAVENPORT	86	PR D33 2519	T.F. Davenport	(TUFTS, ARIZ, FNAL, FSU, NDAM+)
DUKSTRA	86	ZPHY C31 375	H. Dijkstra <i>et al.</i>	(ANIK, BRIS, CERN+)
FRAME	86	NP B276 667	D. Frame <i>et al.</i>	(GLAS)
GOLUBEV	86	SJNP 44 409	V.B. Golubev <i>et al.</i>	(NOVO)
Translated from YAF 44 633.				
ALBRECHT	85D	PL B53B 343	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
GOLUBEV	85	SJNP 41 756	V.B. Golubev <i>et al.</i>	(NOVO)
Translated from YAF 41 1183.				
DRUZHININ	84	PL B44B 136	V.P. Druzhinin <i>et al.</i>	(NOVO)
ARMSTRONG	83B	NP B224 193	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
BARATE	83	PL B121B 449	R. Barate <i>et al.</i>	(SACL, LOIC, SHMP, IND)
KURDADZE	83C	JETPL 38 366	L.M. Kurdadze <i>et al.</i>	(NOVO)
Translated from ZETFP 38 306.				
ARENTON	82	PR D25 2241	M.W. Arenton <i>et al.</i>	(ANL, ILL)
PELLINEN	82	PS 25 599	A. Pellinen, M. Roos	(HEL5)
DAUM	81	PL B100B 439	C. Daum <i>et al.</i>	(AMST, BRIS, CERN, CRAC+)
IVANOV	81	PL B107B 297	P.M. Ivanov <i>et al.</i>	(NOVO)
Also				
VASSERMAN	81	PL B99B 62	I.B. Vasserman <i>et al.</i>	(NOVO)
Also				
SJNP	35	240	L.M. Kurdadze <i>et al.</i>	
Translated from YAF 35 352.				
CORDIER	80	NP B172 13	A. Cordier <i>et al.</i>	(LALO)
CORDIER	79	PL B18B 389	A.D. Cordier <i>et al.</i>	(LALO)
BUKIN	78B	SJNP 27 521	A.D. Bukin <i>et al.</i>	(NOVO)
Translated from YAF 27 985.				

BUKIN	78C	SJNP 27 516	A.D. Bukin <i>et al.</i>	(NOVO)
Translated from YAF 27 976.				
COOPER	78B	NP B146 1	A.M. Cooper <i>et al.</i>	(TATA, CERN, CDEF+)
LOSTY	78	NP B133 38	M.J. Losty <i>et al.</i>	(CERN, AMST, NIUM+)
AKERLOF	77	PRL 39 861	C.W. Akerlof <i>et al.</i>	(FNAL, MICH, PURD)
ANDREWS	77	PRL 38 198	D.E. Andrews <i>et al.</i>	(ROCH)
BALDI	77	PL 68B 381	R. Baldi <i>et al.</i>	(GEVA)
CERRADA	77B	NP B126 241	M. Cerrada <i>et al.</i>	(AMST, CERN, NIUM+)
COHEN	77	PRL 38 269	D. Cohen <i>et al.</i>	(ANL)
LAVER	77	NP B127 43	H. Laven <i>et al.</i>	(AACH3, BERL, CERN, LOIC+)
LYONS	77	NP B125 207	L. Lyons, A.M. Cooper, A.G. Clark	(OXF)
COSME	76	PL 63B 352	G. Cosme <i>et al.</i>	(ORSAY)
KALBFLEISCH	76	PR D13 22	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
PARROUR	76	PL 63B 357	G. Parour <i>et al.</i>	(ORSAY)
PARROUR	76B	PL 63B 362	G. Parour <i>et al.</i>	(ORSAY)
KALBFLEISCH	75	PR D11 987	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
AYRES	74	PRL 32 1463	D.S. Ayres <i>et al.</i>	(ANL)
BESCH	74	NP B70 257	H.J. Besch <i>et al.</i>	(BONN)
COSME	74	PL 48B 155	G. Cosme <i>et al.</i>	(ORSAY)
COSME	74B	PL 48B 159	G. Cosme <i>et al.</i>	(ORSAY)
DEGROOT	74	NP B74 77	A.J. de Groot <i>et al.</i>	(AMST, NIUM)
AUGUSTIN	73	PRL 30 462	J.E. Augustin <i>et al.</i>	(ORSAY)
BALLAM	73	PR D7 3150	J. Ballam <i>et al.</i>	(SLAC, LBL)
BINNIE	73B	PR D8 2789	D.M. Binnie <i>et al.</i>	(LOIC, SHMP)
AGUILAR...	72B	PR D6 29	M. Aguilar-Benitez <i>et al.</i>	(BNL)
ALVENSLEB...	72	PRL 28 66	H. Alvensleben <i>et al.</i>	(MIT, DESY)
BORENSTEIN	72	PR D5 1559	S.R. Borenstein <i>et al.</i>	(BNL, MICH)
COLLEY	72	NP B50 1	D.C. Colley <i>et al.</i>	(BIRM, GLAS)
BALAKIN	71	PL 34B 328	V.E. Balakin <i>et al.</i>	(NOVO)
CHATELUS	71	Thesis LAL 1247	Y. Chatelus	(STRB)
Also				
HAYES	71	PR D4 899	S. Hayes <i>et al.</i>	(CORN)
STOTTLE...	71	Thesis ORO 2504 170	A.R. Stottlemeyer	(UMD)
BIZOT	70	PL 32B 416	J.C. Bizot <i>et al.</i>	(ORSAY)
Also				
EARLES	70	PRL 25 1312	J.P. Perez-y-Jorba	(NEAS)
LINDSEY	66	PR 147 913	J.S. Lindsey, G. Smith	(LRL)
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYR) IJGJC
BADIER	65B	PL 17 337	J. Badier <i>et al.</i>	(EPOL, SACL, AMST)
LINDSEY	65	PRL 15 221	J.S. Lindsey, G.A. Smith	(LRL)
LINDSEY 65 data included in LINDSEY 66.				
SCHLEIN	63	PRL 10 368	P.E. Schlein <i>et al.</i>	(UCLA) IJGJP

 $h_1(1170)$ 

$$I^G(J^{PC}) = 0^-(1^{+-})$$

 $h_1(1170)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>1170±20 OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1168± 4	ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1166± 5±3	<sup>1</sup> ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1190±60	<sup>2</sup> DANKOWY...	81	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
<sup>1</sup> Average and spread of values using 2 variants of the model of BOWLER 75.				
<sup>2</sup> Uses the model of BOWLER 75.				

 $h_1(1170)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>360±40 OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
345± 6	ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
375± 6±34	<sup>3</sup> ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
320±50	<sup>4</sup> DANKOWY...	81	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
<sup>3</sup> Average and spread of values using 2 variants of the model of BOWLER 75.				
<sup>4</sup> Uses the model of BOWLER 75.				

 $h_1(1170)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \rho\pi$	seen

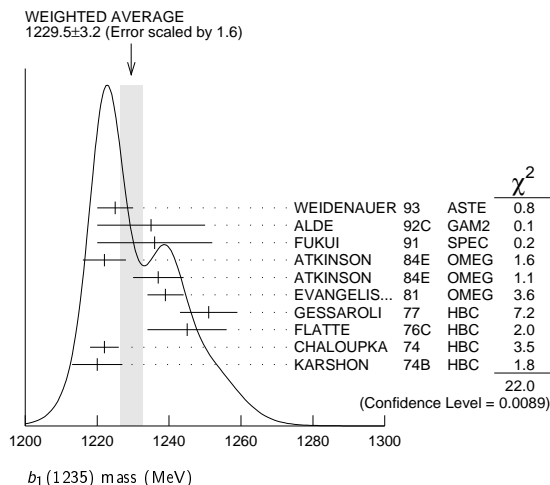
 $h_1(1170)$  BRANCHING RATIOS

$\Gamma(\rho\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
seen	ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
seen	ATKINSON	84	OMEG	$20\text{--}70 \gamma p \rightarrow \pi^+ \pi^- \pi^0 p$
seen	DANKOWY...	81	SPEC	$8 \pi p \rightarrow 3 \pi n$

$$b_1(1235)$$

$$I^G(J^{PC}) = 1^+(1^+ -)$$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1229.5 ± 3.2 OUR AVERAGE</b>		Error includes scale factor of 1.6.			See the ideogram below.
1225 ± 5		WEIDENAUER 93	ASTE		$\bar{p}p \rightarrow 2\pi^+ 2\pi^- \pi^0$
1235 ± 15		ALDE 92C	GAM2		$38,100 \pi^- p \rightarrow \omega \pi^0 n$
1236 ± 16		FUKUI 91	SPEC		$8.95 \pi^- p \rightarrow \omega \pi^0 n$
1222 ± 6		ATKINSON 84E	OMEG ±		$25\text{--}55 \gamma p \rightarrow \omega \pi X$
1237 ± 7		ATKINSON 84E	OMEG 0		$25\text{--}55 \gamma p \rightarrow \omega \pi X$
1239 ± 5		EVANGELIN... 81	OMEG ±		$12 \pi^- p \rightarrow \omega \pi p$
1251 ± 8	450	GESSAROLI 77	HBC	−	$11 \pi^- p \rightarrow \pi^- \omega p$
1245 ± 11	890	FLATTE 76C	HBC	−	$4.2 K^- p \rightarrow \pi^- \omega \Sigma^+$
1222 ± 4	1400	CHALOUKKA 74	HBC	−	$3.9 \pi^- p$
1220 ± 7	600	KARSHON 74B	HBC	+	$4.9 \pi^+ p$
● ● ●	We do not use the following data for averages, fits, limits, etc. ● ● ●				
1190 ± 10		AUGUSTIN 89	DM2 ±		$e^+ e^- \rightarrow 5\pi$
1213 ± 5		ATKINSON 84C	OMEG 0		$20\text{--}70 \gamma p$
1271 ± 11		COLLICK 84	SPEC +		$200 \pi^+ Z \rightarrow Z \pi \omega$

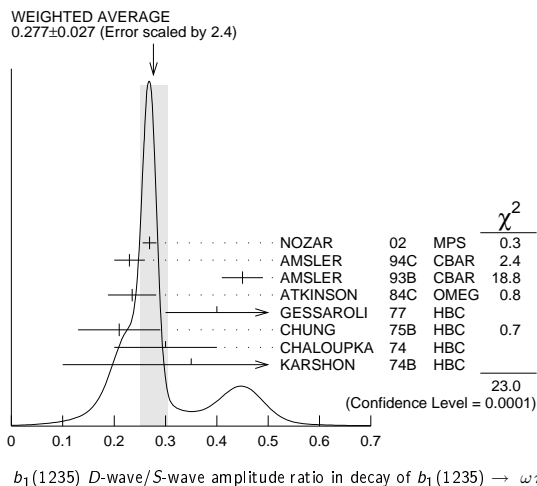


VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b>142± 9 OUR AVERAGE</b>		Error includes scale factor of 1.2.			
113±12		WEIDENAUER 93	ASTE		$\bar{p}p \rightarrow 2\pi^+ 2\pi^- \pi^0$
160±30		ALDE 92c	GAM2		$38,100 \pi^- p \rightarrow \omega \pi^0 n$
151±31		FUKUI 91	SPEC		$8.95 \pi^- p \rightarrow \omega \pi^0 n$
170±15		EVANGELIS... 81	OMEG	−	$12 \pi^- p \rightarrow \omega \pi p$
170±50	225	BALTAY 78B	HBC	+	$15 \pi^+ p \rightarrow p 4\pi$
155±32	450	GESSAROLI 77	HBC	−	$11 \pi^- p \rightarrow \pi^- \omega p$
182±45	890	FLATTE 76c	HBC	−	$4.2 K^- p \rightarrow \pi^- \omega \Sigma^+$
135±20	1400	CHALOUPEKA 74	HBC	−	$3.9 \pi^- p \rightarrow$
156±22	600	KARSHON 74B	HBC	+	$4.9 \pi^+ p$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
210±19		AUGUSTIN 89	DM2	±	$e^+ e^- \rightarrow 5\pi$
231±14		ATKINSON 84c	OMEG	0	$20\text{--}70 \gamma p$
232±29		COLLICK 84	SPEC	+	$200 \pi^+ Z \rightarrow Z \pi \omega$

	Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$	$\omega\pi$ [ $D/S$ amplitude ratio = $0.277 \pm 0.027$ ]	dominant	
$\Gamma_2$	$\pi^\pm\gamma$	$(1.6 \pm 0.4) \times 10^{-3}$	
$\Gamma_3$	$\eta\rho$	seen	
$\Gamma_4$	$\pi^+\pi^+\pi^-\pi^0$	$< 50$ %	84%
$\Gamma_5$	$K^*(892)^\pm K^\mp$	seen	
$\Gamma_6$	$(K\bar{K})^\pm\pi^0$	$< 8$ %	90%
$\Gamma_7$	$K_S^0 K_L^0 \pi^\pm$	$< 6$ %	90%
$\Gamma_8$	$K_S^0 K_S^0 \pi^\pm$	$< 2$ %	90%
$\Gamma_9$	$\phi\pi$	$< 1.5$ %	84%

$\Gamma(\pi^{\pm}\gamma)$					$\Gamma_2$
VALUE (keV)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>230±60</b>	COLLICK	84	SPEC	+	200 $\pi^+\pi^-\gamma \rightarrow Z\pi^+\omega$

VALUE	EVENTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.277 ± 0.027 OUR AVERAGE</b>		Error includes scale factor of 2.4. See the ideogram below.			
0.269 ± 0.009 ± 0.010		NOZAR	02	MP5	- 18 $\pi^- p \rightarrow \omega \pi^- p$
0.23 ± 0.03		AMSLER	94c	CBAR	0.0 $\overline{p} p \rightarrow \omega \eta \pi^0$
0.45 ± 0.04		AMSLER	93b	CBAR	0.0 $\overline{p} p \rightarrow \omega \pi^0 \pi^0$
0.235 ± 0.047		AT KINSON	84c	OMEG	20-70 $\gamma p$
0.4 $\begin{smallmatrix} +0.1 \\ -0.1 \end{smallmatrix}$		GESSAROLI	77	HBC	- 11 $\pi^- p \rightarrow \pi^- \omega p$
0.21 ± 0.08		CHUNG	75b	HBC	+ 7.1 $\pi^+ p$
0.3 ± 0.1		CHALOUPEKA	74	HBC	- 3.9-7.5 $\pi^- p$
0.35 ± 0.25	600	KARSHON	74b	HBC	+ 4.9 $\pi^+ p$



<u>VALUE (°)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>10.5±2.4±3.9</b>	NOZAR	02	MPS	— 18 $\pi^- p \rightarrow \omega \pi^- p$

$\Gamma(\eta\rho)/\Gamma(\omega\pi)$				$\Gamma_3/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.10</b>	ATKINSON	84D	OMEG	20-70 $\gamma$ p

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\omega\pi)$	$\Gamma_4/\Gamma_1$			
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;0.5</b>	ABOLINS	63	HBC	+ 3.5 $\pi^+ p$

$\Gamma(K^*(892)^\pm K^\mp)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	<sup>1</sup> ABLIKIM	10E BES2	$J/\psi \rightarrow K^\pm K^0 \pi^\mp \pi^0$

<sup>1</sup> From a fit including ten additional resonances and energy-independent Breit-Wigner width.

$\Gamma((K\bar{K})^{\pm}\pi^0)/\Gamma(\omega\pi)$					$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<0.08	90	BAITAY 67	HRC	+	0.0 $\overline{p}n$

$\Gamma(K_S^0 K_L^0 \pi^\pm)/\Gamma(\omega\pi)$					$\Gamma_7/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.06	90	RAITAY	67	HRC	+	0.0 $\overline{p}n$

$\Gamma(K_S^0 K_S^0 \pi^\pm)/\Gamma(\omega\pi)$					$\Gamma_8/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.02	90	RAITAY	67	HRC	+	0.0 $\overline{20}$

See key on page 601

## Meson Particle Listings

 $b_1(1235)$ ,  $a_1(1260)$ 

$\Gamma(\phi\pi)/\Gamma(\omega\pi)$					$\Gamma_9/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<0.004	95	VIKTOROV	96	SPEC	0 32.5 $\pi^- p \rightarrow K^+ K^- \pi^0 n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.04	95	BIZZARRI	69	HBC	$\pm$ 0.0 $\overline{p} p$
<0.015		DAHL	67	HBC	1.6–4.2 $\pi^- p$

 $b_1(1235)$  REFERENCES

ABLIKIM	10E	PL B693 88	M. Ablikim <i>et al.</i>	(BES II Collab.)
NOZAR	02	PL B541 35	M. Nozar <i>et al.</i>	
VIKTOROV	96	PAN 59 1104	V.A. Viktorov <i>et al.</i>	(SERP)
Translated from YAF 59 1239				
AMSLER	94C	PL B327 425	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	93B	PL B311 362	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
WEIDENAUER	93	ZPHY C59 387	P. Weidenauer <i>et al.</i>	(ASTERIX Collab.)
ALDE	92C	ZPHY C54 553	D.M. Alde <i>et al.</i>	(BELG, SERP, KEK, LANL+)
FUKUI	91	PL B257 241	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
AUGUSTIN	89	NP B320 1	J.E. Augustin, G. Cosme	(DM2 Collab.)
ATKINSON	84C	NP B243 1	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
ATKINSON	84D	NP B242 269	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
ATKINSON	84E	PL 130B 459	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
COLLICK	84	PRL 53 2374	B. Collick <i>et al.</i>	(MINN, ROCH, FNAL)
EVANGELIS...	81	NP B178 197	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
BALTAY	76B	PR D17 62	C. Baltay <i>et al.</i>	(COLU, BING)
GESSAROLI	77	NP B126 382	R. Gessaroli <i>et al.</i>	(BGNA, FIRZ, GENO+)
FLATTE	76C	PL 64B 225	S.M. Flatte <i>et al.</i>	(CERN, AMST, NIJN+)
CHUNG	75B	PR D11 2426	S.U. Chung <i>et al.</i>	(BNL, LBL, UCSC)
CHALOUPKA	74	PL 51B 407	V. Chaloupka <i>et al.</i>	(CERN)
KARSHON	74B	PR D10 3608	U. Karshon <i>et al.</i>	(REHO)JP
BIZZARRI	69	NP B14 169	R. Bizzarri <i>et al.</i>	(CERN, CDEF)
BALTAY	67	PRL 18 93	C. Baltay <i>et al.</i>	(COLU)
DAHL	67	PR 163 1377	O.I. Dahl <i>et al.</i>	(LRL)
ABOLINS	63	PRL 11 381	M.A. Abolins <i>et al.</i>	(UCSD)

 $a_1(1260)$ 

$$I^G(J^{PC}) = 1^-(1^{++})$$

See also our review under the  $a_1(1260)$  in PDG 06, Journal of Physics **G33** 1 (2006).

 $a_1(1260)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1230±40 OUR ESTIMATE</b>				
<b>1255±6<sup>+7</sup><sub>-17</sub></b>	420k	ALEKSEEV	10	COMP 190 $\pi^- P_b \rightarrow \pi^- \pi^- \pi^+ P_b'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1243±12±20		<sup>1</sup> AUBERT	07AU BABR	10.6 $e^+ e^- \rightarrow \rho^0 \rho^\pm \pi^\mp \gamma$
1230–1270	6360	<sup>2</sup> LINK	07A FOCs	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
1203±3		<sup>3</sup> GOMEZ-DUM..04	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
1330±24	90k	SALVINI	04 OBLX	$\overline{p} p \rightarrow 2\pi^+ 2\pi^-$
1331±10±3	37k	<sup>4</sup> ASNER	00 CLE2	10.6 $e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
1255±7±6	5904	<sup>5</sup> ABREU	98G DLPH	$e^+ e^-$
1207±5±8	5904	<sup>6</sup> ABREU	98G DLPH	$e^+ e^-$
1196±4±5	5904	<sup>7,8</sup> ABREU	98G DLPH	$e^+ e^-$
1240±10		BARBERIS	98B	450 $pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
1262±9±7		<sup>5,9</sup> ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
1210±7±2		<sup>6,9</sup> ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
1211±7 <sup>+50</sup> <sub>-0</sub>		<sup>6</sup> ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1121±8		<sup>10</sup> ANDO	92 SPEC	8 $\pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1242±37		<sup>11</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260±14		<sup>12</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1250±9		<sup>13</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1208±15		ARMSTRONG	90 OMEG	300.0 $pp \rightarrow p p \pi^+ \pi^- \pi^0$
1220±15		<sup>14</sup> ISGUR	89 RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260±25		<sup>15</sup> BOWLER	88 RVUE	
1166±18±11		BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1164±41±23		BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
1250±40		<sup>14</sup> TORNQVIST	87 RVUE	
1046±11		ALBRECHT	86B ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1056±20±15		RUCKSTUHL	86 DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1194±14±10		SCHMIDKE	86 MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1255±23		BELLINI	85 SPEC	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1240±80		<sup>16</sup> DANKOWY...	81 SPEC	8.45 $\pi^- p \rightarrow n 3\pi$
1280±30		<sup>16</sup> DAUM	81B CNTR	63,94 $\pi^- p \rightarrow p 3\pi$
1041±13		<sup>17</sup> GAVILLET	77 HBC	4.2 $K^- p \rightarrow \Sigma 3\pi$

<sup>1</sup> The  $\rho^\pm \pi^\mp$  state can be also due to the  $\pi(1300)$ .

<sup>2</sup> Using the Breit-Wigner parameterization; strong correlation between mass and width.

<sup>3</sup> Using the data of BARATE 98R.

<sup>4</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\overline{K}^*(892)$  threshold.

<sup>5</sup> Uses the model of KUHN 90.

<sup>6</sup> Uses the model of ISGUR 89.

<sup>7</sup> Includes the effect of a possible  $a_1'$  state.

<sup>8</sup> Uses the model of FEINDT 90.

<sup>9</sup> Supersedes AKERS 95P.

<sup>10</sup> Average and spread of values using 2 variants of the model of BOWLER 75.

<sup>11</sup> Reanalysis of RUCKSTUHL 86.

<sup>12</sup> Reanalysis of SCHMIDKE 86.

<sup>13</sup> Reanalysis of ALBRECHT 86B.

<sup>14</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.

<sup>15</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.

<sup>16</sup> Uses the model of BOWLER 75.

<sup>17</sup> Produced in  $K^-$  backward scattering.

 $a_1(1260)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>250 to 600 OUR ESTIMATE</b>				
<b>367±9<sup>+28</sup><sub>-25</sub></b>	420k	ALEKSEEV	10	COMP 190 $\pi^- P_b \rightarrow \pi^- \pi^- \pi^+ P_b'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
410±31±30		<sup>18</sup> AUBERT	07AU BABR	10.6 $e^+ e^- \rightarrow \rho^0 \rho^\pm \pi^\mp \gamma$
520–680	6360	<sup>19</sup> LINK	07A FOCs	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
480±20		<sup>20</sup> GOMEZ-DUM..04	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
580±41	90k	SALVINI	04 OBLX	$\overline{p} p \rightarrow 2\pi^+ 2\pi^-$
460±85	205	<sup>21</sup> DRUTSKOY	02 BELL	$B \rightarrow D^{(*)} K^- K^* 0$
814±36±13	37k	<sup>22</sup> ASNER	00 CLE2	10.6 $e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
450±50	22k	<sup>23</sup> AKHMETSHIN 99E	CMD2	1.05–1.38 $e^+ e^- \rightarrow$ $\pi^+ \pi^- \pi^0 \pi^0$
570±10		<sup>24</sup> BONDAR	99 RVUE	$e^+ e^- \rightarrow 4\pi, \tau \rightarrow 3\pi \nu_\tau$
587±27±21	5904	<sup>25</sup> ABREU	98G DLPH	$e^+ e^-$
478±3±15	5904	<sup>26</sup> ABREU	98G DLPH	$e^+ e^-$
425±14±8	5904	<sup>27,28</sup> ABREU	98G DLPH	$e^+ e^-$
400±35		BARBERIS	98B	450 $pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
621±32±58		<sup>25,29</sup> ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
457±15±17		<sup>26,29</sup> ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
446±21 <sup>+140</sup> <sub>-0</sub>		<sup>26</sup> ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
239±11		ANDO	92 SPEC	8 $\pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
266±13±4		<sup>30</sup> ANDO	92 SPEC	8 $\pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
465 <sup>+228</sup> <sub>-143</sub>		<sup>31</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
298±40 <sub>-34</sub>		<sup>32</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
488±32		<sup>33</sup> IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
430±50		ARMSTRONG	90 OMEG	300.0 $pp \rightarrow p p \pi^+ \pi^- \pi^0$
420±40		<sup>34</sup> ISGUR	89 RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
396±43		<sup>35</sup> BOWLER	88 RVUE	
405±75±25		BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
419±108±57		BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
521±27		ALBRECHT	86B ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
476 <sup>+132</sup> <sub>-120</sub> ±54		RUCKSTUHL	86 DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
462±56±30		SCHMIDKE	86 MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
292±40		BELLINI	85 SPEC	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
380±100		<sup>36</sup> DANKOWY...	81 SPEC	8.45 $\pi^- p \rightarrow n 3\pi$
300±50		<sup>36</sup> DAUM	81B CNTR	63,94 $\pi^- p \rightarrow p 3\pi$
230±50		<sup>37</sup> GAVILLET	77 HBC	4.2 $K^- p \rightarrow \Sigma 3\pi$

<sup>18</sup> The  $\rho^\pm \pi^\mp$  state can be also due to the  $\pi(1300)$ .

<sup>19</sup> Using the Breit-Wigner parameterization; strong correlation between mass and width.

<sup>20</sup> Using the data of BARATE 98R.

<sup>21</sup> From a fit of the  $K^- K^* 0$  distribution assuming  $m_{a_1} = 1230$  MeV and purely resonant production of the  $K^- K^* 0$  system.

<sup>22</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\overline{K}^*(892)$  threshold.

<sup>23</sup> Using the  $a_1(1260)$  mass of 1230 MeV.

<sup>24</sup> From AKHMETSHIN 99E and ASNER 00 data using the  $a_1(1260)$  mass of 1230 MeV.

<sup>25</sup> Uses the model of KUHN 90.

<sup>26</sup> Uses the model of ISGUR 89.

<sup>27</sup> Includes the effect of a possible  $a_1'$  state.

<sup>28</sup> Uses the model of FEINDT 90.

<sup>29</sup> Supersedes AKERS 95P.

<sup>30</sup> Average and spread of values using 2 variants of the model of BOWLER 75.

<sup>31</sup> Reanalysis of RUCKSTUHL 86.

<sup>32</sup> Reanalysis of SCHMIDKE 86.

<sup>33</sup> Reanalysis of ALBRECHT 86B.

<sup>34</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.

<sup>35</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.

<sup>36</sup> Uses the model of BOWLER 75.

<sup>37</sup> Produced in  $K^-$  backward scattering.

 $a_1(1260)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi^+ \pi^- \pi^0$	
$\Gamma_2$ $\pi^0 \pi^0 \pi^0$	
$\Gamma_3$ $(\rho\pi)S$ -wave	seen
$\Gamma_4$ $(\rho\pi)D$ -wave	seen
$\Gamma_5$ $(\rho(1450)\pi)S$ -wave	seen
$\Gamma_6$ $(\rho(1450)\pi)D$ -wave	seen

# Meson Particle Listings

## $a_1(1260)$

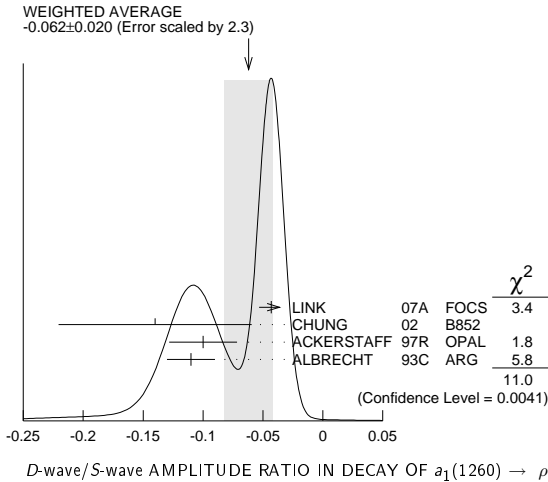
$\Gamma_7$	$\sigma\pi$	seen
$\Gamma_8$	$f_0(980)\pi$	not seen
$\Gamma_9$	$f_0(1370)\pi$	seen
$\Gamma_{10}$	$f_2(1270)\pi$	seen
$\Gamma_{11}$	$K\bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{12}$	$\pi\gamma$	seen

### $a_1(1260)$ PARTIAL WIDTHS

$\Gamma(\pi\gamma)$	VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12}$
<b><math>640 \pm 246</math></b>		ZIELINSKI	84c	SPEC	$200\pi^+\text{Z} \rightarrow \text{Z}\pi$

### D-wave/S-wave AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.062 \pm 0.020</math> OUR AVERAGE</b>	Error includes scale factor of 2.3. See the ideogram below.		
$-0.043 \pm 0.009 \pm 0.005$	LINK	07A	FOCS $D^0 \rightarrow \pi^-\pi^+\pi^-\pi^+$
$-0.14 \pm 0.04 \pm 0.07$	<sup>38</sup> CHUNG	02	B852 $18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$
$-0.10 \pm 0.02 \pm 0.02$	<sup>39,40</sup> ACKERSTAFF	97R	OPAL $E_{\text{cm}}^{\text{pe}} = 88\text{--}94, \tau \rightarrow 3\pi\nu$
$-0.11 \pm 0.02$	<sup>39</sup> ALBRECHT	93c	ARG $\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
<sup>38</sup> Deck-type background not subtracted.			
<sup>39</sup> Uses the model of ISGUR 89.			
<sup>40</sup> Supersedes AKERS 95P.			



### $a_1(1260)$ BRANCHING RATIOS

$\Gamma((\rho\pi)\text{s-wave})/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$60.19$		37k	<sup>41</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma((\rho\pi)\text{D-wave})/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$1.30 \pm 0.60 \pm 0.22$		37k	<sup>41</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma((\rho(1450)\pi)\text{s-wave})/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.56 \pm 0.84 \pm 0.32$		37k	<sup>41,42</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma((\rho(1450)\pi)\text{D-wave})/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$2.04 \pm 1.20 \pm 0.28$		37k	<sup>41,42</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma(\sigma\pi)/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
seen			CHUNG	02	B852 $18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$	
$18.76 \pm 4.29 \pm 1.48$		37k	<sup>41,43</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma(f_0(980)\pi)/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
not seen		37k	ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma(f_0(1370)\pi)/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$7.40 \pm 2.71 \pm 1.26$		37k	<sup>41,44</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma(f_2(1270)\pi)/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$1.19 \pm 0.49 \pm 0.17$		37k	<sup>41,45</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	

$\Gamma(K\bar{K}^*(892) + \text{c.c.})/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$2.2 \pm 0.5$		2255	<sup>46</sup> COAN	04	CLEO $\tau^- \rightarrow K^-\pi^-K^+\nu_\tau$	
$8 \text{ to } 15$		205	<sup>47</sup> DRUTSKOY	02	BELL $B \rightarrow D^{(*)}K\bar{K}^*0$	
$3.3 \pm 0.5 \pm 0.1$		37k	<sup>48</sup> ASNER	00	CLE2 $10.6e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$	
$2.6 \pm 0.3$			<sup>49</sup> BARATE	99R	ALEP $\tau \rightarrow K\bar{K}\pi\nu_\tau$	

$\Gamma(\sigma\pi)/\Gamma((\rho\pi)\text{s-wave})$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma_3$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.06 \pm 0.05$		90k	SALVINI	04	OBLX $\bar{p}p \rightarrow 2\pi^+2\pi^-$	
$\sim 0.3$		28k	AKHMETSHIN 99E	CMD2	$1.05\text{--}1.38e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$	
$0.003 \pm 0.003$			<sup>50</sup> LONGACRE	82	RVUE	

$\Gamma(\pi^0\pi^0\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$	VALUE	CL%	DOCUMENT ID	COMMENT	$\Gamma_2/\Gamma_1$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<0.008$		90	<sup>51</sup> BARBERIS	01 450 $pp \rightarrow p_f 3\pi^0 p_s$	

- <sup>41</sup> From a fit to the Dalitz plot.  
<sup>42</sup> Assuming for  $\rho(1450)$  mass and width of 1370 and 386 MeV respectively.  
<sup>43</sup> Assuming for  $\sigma$  mass and width of 860 and 880 MeV respectively.  
<sup>44</sup> Assuming for  $f_0(1370)$  mass and width of 1186 and 350 MeV respectively.  
<sup>45</sup> Assuming for  $f_2(1270)$  mass and width of 1275 and 185 MeV respectively.  
<sup>46</sup> Using structure functions from KUNN 92 and DECKER 93A and B( $\tau^- \rightarrow K^-\pi^-K^+\nu_\tau = (0.155 \pm 0.006 \pm 0.009)\%$  from BRIERE 03).  
<sup>47</sup> From a comparison to ALAM 94 assuming purely resonant production of the  $K^-\bar{K}^*0$  system.  
<sup>48</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*(892)$  threshold.  
<sup>49</sup> Assuming  $a_1(1260)$  dominance and taking B( $\tau \rightarrow a_1(1260)\nu_\tau$ ) from BUSKULIC 96.  
<sup>50</sup> Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from GAVIL-LET 77, DAUM 80, and DANKOWYCH 81.  
<sup>51</sup> Inconsistent with observations of  $\sigma\pi$ ,  $f_0(1370)\pi$ , and  $f_2(1270)\pi$  decay modes.

### $a_1(1260)$ REFERENCES

ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i>	(COMPASS Collab.)
AUBERT	07AU	PR D76 092005	B. Aubert <i>et al.</i>	(BABAR Collab.)
LINK	07A	PR D75 052003	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
COAN	04	PRL 92 232001	T.E. Coan <i>et al.</i>	(CLEO Collab.)
GOMEZ-DUMM	04	PR D69 073002	D. Gomez Dumm, A. Pich, J. Portoles	
SALVINI	04	EPJ C35 21	P. Salvini <i>et al.</i>	(OBELIX Collab.)
BRIERE	03	PRL 90 181802	R. A. Briere <i>et al.</i>	(CLEO Collab.)
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
DRUTSKOY	02	PL B542 171	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
BARBERIS	01	PL B507 14	D. Barberis <i>et al.</i>	
ASNER	00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
AKHMETSHIN	99E	PL B466 392	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BARATE	99R	EPJ C11 599	R. Barate <i>et al.</i>	(ALEPH Collab.)
BONDAR	99	PL B466 403	A.E. Bondar <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ABREU	98G	PL B426 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BARATE	98R	EPJ C4 409	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARBERIS	98B	PL B422 339	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ACKERSTAFF	97R	ZPHY C75 593	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BUSKULIC	96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
ALAM	94	PR D50 43	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DECKER	93A	ZPHY C58 445	R. Decker <i>et al.</i>	
ANDO	92	PL B291 496	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
KUNN	92	ZPHY C56 661	J.H. Kuhn, E. Mirkes	
IVANOV	91	ZPHY C49 563	Y.P. Ivanov, A.A. Osipov, M.K. Volkov	(JINR)
ARMSTRONG	90	ZPHY C48 213	T.A. Armstrong, M. Benayoun, W. Beusch	(WA76 Coll.)
FEINDT	90	ZPHY C48 681	M. Feindt	(HAMB)
KUNN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MFM)
ISGUR	89	PR D39 1357	N. Isgur, C. Morningstar, C. Reader	(TNTO)
BOWLER	88	PL B209 99	M.G. Bowler	(OXF)
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
TORNQVIST	87	ZPHY C36 695	N.A. Tornqvist	(HELS)
ALBRECHT	86B	ZPHY C33 7	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
RUCKSTUHL	86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)
SCHMIDKE	86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)
BELLINI	85	SJNP 41 781	D. Bellini <i>et al.</i>	

Translated from YAF 41 1223.

See key on page 601

## Meson Particle Listings

 $a_1(1260)$ ,  $f_2(1270)$ 

ZIELINSKI	84C	PRL 52 1195	M. Zielinski <i>et al.</i>	(ROCH, MINN, FNAL)
LONGACRE	82	PR D26 82	R.S. Longacre	(BNL)
DANKOWY...	81	PRL 46 580	J.A. Dankowycz <i>et al.</i>	(TNTO, BNL, CARL+)
DAUM	81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
DAUM	80	PL B9B 281	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
GAUILLET	77	PL 69B 119	P. Gavillet <i>et al.</i>	(AMST, CERN, NUM+)
BOWLER	75	NP B97 227	M.G. Bowler <i>et al.</i>	(OXFTP, DARE)

 $f_2(1270)$ 

$$I^G(J^{PC}) = 0^+(2^{++})$$

 $f_2(1270)$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1275.5 ± 0.8 OUR AVERAGE</b>				
1275.8 ± 1.0 ± 0.4		<sup>1</sup> BOGOLYUB... 13	SPEC	$7\pi^+(K^+,p)A \rightarrow n\gamma + X$
1262 ± 1 ± 8		ABLIKIM 06v	BES2	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^+\pi^-$
1275 ± 15		ABLIKIM 05	BES2	$J/\psi \rightarrow \phi\pi^+\pi^-$
1283 ± 5		ALDE 98	GAM4	$100\pi^-p \rightarrow \pi^0\pi^0n$
1278 ± 5		<sup>2</sup> BERTIN 97c	OBLX	$0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
1272 ± 8	200k	PROKOSHKIN 94	GAM2	$38\pi^-p \rightarrow \pi^0\pi^0n$
1269.7 ± 5.2	5730	AUGUSTIN 89	DM2	$e^+e^- \rightarrow 5\pi$
1283 ± 8	400	<sup>3</sup> ALDE 87	GAM4	$100\pi^-p \rightarrow 4\pi^0n$
1274 ± 5		<sup>3</sup> AUGUSTIN 87	DM2	$J/\psi \rightarrow \gamma\pi^+\pi^-$
1283 ± 6		<sup>4</sup> LONGACRE 86	MPS	$22\pi^-p \rightarrow n2K_S^0$
1276 ± 7		COURAU 84	DLCO	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
1273.3 ± 2.3		<sup>5</sup> CHABAUD 83	ASPK	$17\pi^-p$ polarized
1280 ± 4		<sup>6</sup> CASON 82	STRC	$8\pi^+p \rightarrow \Delta^{++}\pi^0\pi^0$
1281 ± 7	11600	GIDAL 81	MRK2	$J/\psi$ decay
1282 ± 5		<sup>7</sup> CORDEN 79	OMEG	$12-15\pi^-p \rightarrow n2\pi$
1269 ± 4	10k	APEL 75	NICE	$40\pi^-p \rightarrow n2\pi^0$
1272 ± 4	4600	ENGLER 74	DBC	$6\pi^+n \rightarrow \pi^+\pi^-p$
1277 ± 4	5300	FLATTE 71	HBC	$7.0\pi^+p$
1273 ± 8		<sup>3</sup> STUNTEBECK 70	HBC	$8\pi^-p, 5.4\pi^+d$
1265 ± 8		BOESEBECK 68	HBC	$8\pi^+p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1259 ± 4 ± 4	1.7k	<sup>8,9</sup> DOBBS 15		$J/\psi \rightarrow \gamma\pi^+\pi^-$
1267 ± 4 ± 3	1.5k	<sup>8,9</sup> DOBBS 15		$\psi(2S) \rightarrow \gamma\pi^+\pi^-$
1270 ± 8		<sup>10</sup> ANISOVICH 09	RVUE	$0.0\bar{p}p, \pi N$
1277 ± 6	870	<sup>11</sup> SCHEGELSKY 06a	RVUE	$\gamma\gamma \rightarrow K_S^0 K_S^0$
1251 ± 10		TIKHOMIROV 03	SPEC	$40.0\pi^-C \rightarrow K_S^0 K_S^0 K_L^0 X$
1260 ± 10		<sup>12</sup> ALDE 97	GAM2	$450pp \rightarrow pp\pi^0\pi^0$
1278 ± 6		<sup>12</sup> GRYGOREV 96	SPEC	$40\pi^-N \rightarrow K_S^0 K_S^0 X$
1262 ± 11		AGUILAR-... 91	EHS	$400pp$
1275 ± 10		AKER 91	CBAR	$0.0\bar{p}p \rightarrow 3\pi^0$
1220 ± 10		BREAKSTONE 90	SFM	$pp \rightarrow pp\pi^+\pi^-$
1288 ± 12		ABACHI 86b	HRS	$e^+e^- \rightarrow \pi^+\pi^-X$
1284 ± 30	3k	BINON 83	GAM2	$38\pi^-p \rightarrow n2\eta$
1280 ± 20	3k	APEL 82	CNTR	$25\pi^-p \rightarrow n2\pi^0$
1284 ± 10	16000	DEUTSCH... 76	HBC	$16\pi^+p$
1258 ± 10	600	TAKAHASHI 72	HBC	$8\pi^-p \rightarrow n2\pi$
1275 ± 13		ARMENISE 70	HBC	$9\pi^+n \rightarrow p\pi^+\pi^-$
1261 ± 5	1960	<sup>3</sup> ARMENISE 68	DBC	$5.1\pi^+n \rightarrow p\pi^+MM^-$
1270 ± 10	360	<sup>3</sup> ARMENISE 68	DBC	$5.1\pi^+n \rightarrow p\pi^0MM$
1268 ± 6		<sup>13</sup> JOHNSON 68	HBC	$3.7-4.2\pi^-p$

<sup>1</sup> Averaged over six nuclear targets, no statistically significant dependence on target nucleus observed.

<sup>2</sup> T-matrix pole.

<sup>3</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>4</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles.

<sup>5</sup> From an energy-independent partial-wave analysis.

<sup>6</sup> From an amplitude analysis of the reaction  $\pi^+\pi^- \rightarrow 2\pi^0$ .

<sup>7</sup> From an amplitude analysis of  $\pi^+\pi^- \rightarrow \pi^+\pi^-$  scattering data.

<sup>8</sup> Using CLEO-c data but not authored by the CLEO Collaboration.

<sup>9</sup> From a fit to a Breit-Wigner line shape with fixed  $\Gamma = 185$  MeV.

<sup>10</sup> 4-poles, 5-channel K matrix fit.

<sup>11</sup> From analysis of L3 data at 91 and 183–209 GeV.

<sup>12</sup> Systematic uncertainties not estimated.

<sup>13</sup> JOHNSON 68 includes BONDAR 63, LEE 64, DERADO 65, EISNER 67.

 $f_2(1270)$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>186.7 ± 2.2 OUR FIT</b>				Error includes scale factor of 1.4.
<b>185.9 ± 2.8 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
190.3 ± 1.9 ± 1.8		<sup>1</sup> BOGOLYUB... 13	SPEC	$7\pi^+(K^+,p)A \rightarrow n\gamma + X$
175 ± 6 ± 10		ABLIKIM 06v	BES2	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^+\pi^-$
190 ± 20		ABLIKIM 05	BES2	$J/\psi \rightarrow \phi\pi^+\pi^-$

171 ± 10		ALDE 98	GAM4	$100\pi^-p \rightarrow \pi^0\pi^0n$
204 ± 20		<sup>2</sup> BERTIN 97c	OBLX	$0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
192 ± 5	200k	PROKOSHKIN 94	GAM2	$38\pi^-p \rightarrow \pi^0\pi^0n$
180 ± 24		AGUILAR-... 91	EHS	$400pp$
169 ± 9	5730	<sup>3</sup> AUGUSTIN 89	DM2	$e^+e^- \rightarrow 5\pi$
150 ± 30	400	<sup>3</sup> ALDE 87	GAM4	$100\pi^-p \rightarrow 4\pi^0n$
186 ± 9 ± 2		<sup>4</sup> LONGACRE 86	MPS	$22\pi^-p \rightarrow n2K_S^0$
179.2 ± 6.9 ± 6.6		<sup>5</sup> CHABAUD 83	ASPK	$17\pi^-p$ polarized
160 ± 11		DENNEY 83	LASS	$10\pi^+N$
196 ± 10	3k	APEL 82	CNTR	$25\pi^-p \rightarrow n2\pi^0$
152 ± 9		<sup>6</sup> CASON 82	STRC	$8\pi^+p \rightarrow \Delta^{++}\pi^0\pi^0$
186 ± 27	11600	GIDAL 81	MRK2	$J/\psi$ decay
216 ± 13		<sup>7</sup> CORDEN 79	OMEG	$12-15\pi^-p \rightarrow n2\pi$
190 ± 10	10k	APEL 75	NICE	$40\pi^-p \rightarrow n2\pi^0$
192 ± 16	4600	ENGLER 74	DBC	$6\pi^+n \rightarrow \pi^+\pi^-p$
183 ± 15	5300	FLATTE 71	HBC	$7\pi^+p \rightarrow \Delta^{++}f_2$
196 ± 30		<sup>3</sup> STUNTEBECK 70	HBC	$8\pi^-p, 5.4\pi^+d$
216 ± 20	1960	<sup>3</sup> ARMENISE 68	DBC	$5.1\pi^+n \rightarrow p\pi^+MM^-$
128 ± 27		<sup>3</sup> BOESEBECK 68	HBC	$8\pi^+p$
176 ± 21		<sup>3,8</sup> JOHNSON 68	HBC	$3.7-4.2\pi^-p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
194 ± 36		<sup>9</sup> ANISOVICH 09	RVUE	$0.0\bar{p}p, \pi N$
195 ± 15	870	<sup>10</sup> SCHEGELSKY 06a	RVUE	$\gamma\gamma \rightarrow K_S^0 K_S^0$
121 ± 26		TIKHOMIROV 03	SPEC	$40.0\pi^-C \rightarrow K_S^0 K_S^0 K_L^0 X$
187 ± 20		<sup>11</sup> ALDE 97	GAM2	$450pp \rightarrow pp\pi^0\pi^0$
184 ± 10		<sup>11</sup> GRYGOREV 96	SPEC	$40\pi^-N \rightarrow K_S^0 K_S^0 X$
200 ± 10		AKER 91	CBAR	$0.0\bar{p}p \rightarrow 3\pi^0$
240 ± 40	3k	BINON 83	GAM2	$38\pi^-p \rightarrow n2\eta$
187 ± 30	650	<sup>3</sup> ANTIPOV 77	CIBS	$25\pi^-p \rightarrow p3\pi$
225 ± 38	16000	DEUTSCH... 76	HBC	$16\pi^+p$
166 ± 28	600	<sup>3</sup> TAKAHASHI 72	HBC	$8\pi^-p \rightarrow n2\pi$
173 ± 53		<sup>3</sup> ARMENISE 70	HBC	$9\pi^+n \rightarrow p\pi^+\pi^-$

<sup>1</sup> Averaged over six nuclear targets, no statistically significant dependence on target nucleus observed.

<sup>2</sup> T-matrix pole.

<sup>3</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>4</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles.

<sup>5</sup> From an energy-independent partial-wave analysis.

<sup>6</sup> From an amplitude analysis of the reaction  $\pi^+\pi^- \rightarrow 2\pi^0$ .

<sup>7</sup> From an amplitude analysis of  $\pi^+\pi^- \rightarrow \pi^+\pi^-$  scattering data.

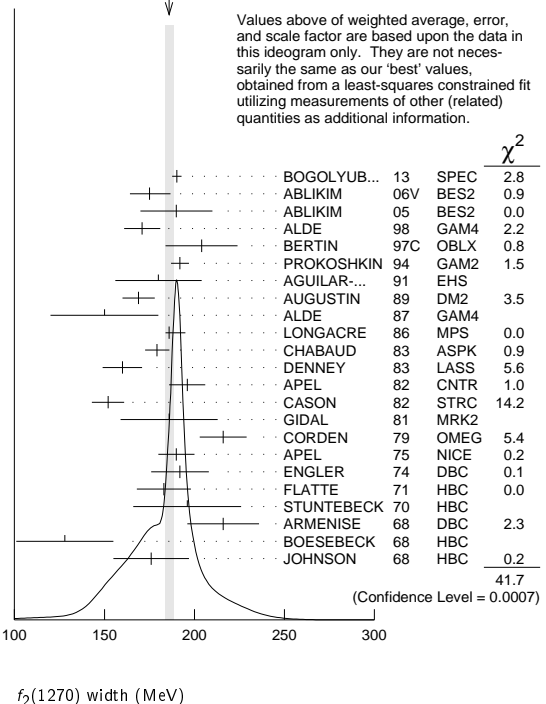
<sup>8</sup> JOHNSON 68 includes BONDAR 63, LEE 64, DERADO 65, EISNER 67.

<sup>9</sup> 4-poles, 5-channel K matrix fit.

<sup>10</sup> From analysis of L3 data at 91 and 183–209 GeV.

<sup>11</sup> Systematic uncertainties not estimated.

WEIGHTED AVERAGE  
185.9±2.8-2.1 (Error scaled by 1.6)



# Meson Particle Listings

## $f_2(1270)$

$f_2(1270)$ DECAY MODES			
Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level	
$\Gamma_1 \quad \pi\pi$	$(84.2 \pm^{+2.9}_{-0.9})\%$	S=1.1	
$\Gamma_2 \quad \pi^+\pi^-2\pi^0$	$(7.7 \pm^{+1.1}_{-3.2})\%$	S=1.2	
$\Gamma_3 \quad K\bar{K}$	$(4.6 \pm^{+0.5}_{-0.4})\%$	S=2.7	
$\Gamma_4 \quad 2\pi^+2\pi^-$	$(2.8 \pm 0.4)\%$	S=1.2	
$\Gamma_5 \quad \eta\eta$	$(4.0 \pm 0.8) \times 10^{-3}$	S=2.1	
$\Gamma_6 \quad 4\pi^0$	$(3.0 \pm 1.0) \times 10^{-3}$		
$\Gamma_7 \quad \gamma\gamma$	$(1.42 \pm 0.24) \times 10^{-5}$	S=1.4	
$\Gamma_8 \quad \eta\pi\pi$	$< 8 \times 10^{-3}$	CL=95%	
$\Gamma_9 \quad K^0\bar{K}^0\pi^+\pi^- + \text{c.c.}$	$< 3.4 \times 10^{-3}$	CL=95%	
$\Gamma_{10} \quad e^+e^-$	$< 6 \times 10^{-10}$	CL=90%	

### CONSTRAINED FIT INFORMATION

An overall fit to the total width, 4 partial widths, a combination of partial widths obtained from integrated cross sections, and 6 branching ratios uses 45 measurements and one constraint to determine 8 parameters. The overall fit has a  $\chi^2 = 83.0$  for 38 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−90						
$x_3$	10	−39					
$x_4$	10	−38	1				
$x_5$	1	−6	0	0			
$x_6$	0	−7	0	0	0		
$x_7$	3	1	−15	0	0	0	
$\Gamma$	−71	65	−10	−7	−1	0	−6
	$x_1$	$x_2$	$x_3$	$x_4$	$x_5$	$x_6$	$x_7$

Mode	Rate (MeV)	Scale factor
$\Gamma_1 \quad \pi\pi$	$157.2 \pm^{+4.0}_{-1.1}$	
$\Gamma_2 \quad \pi^+\pi^-2\pi^0$	$14.4 \pm^{+2.1}_{-6.0}$	1.2
$\Gamma_3 \quad K\bar{K}$	$8.5 \pm 0.8$	2.8
$\Gamma_4 \quad 2\pi^+2\pi^-$	$5.2 \pm 0.7$	1.2
$\Gamma_5 \quad \eta\eta$	$0.75 \pm 0.14$	2.1
$\Gamma_6 \quad 4\pi^0$	$0.56 \pm 0.19$	
$\Gamma_7 \quad \gamma\gamma$	$0.0026 \pm 0.0005$	1.4

### $f_2(1270)$ PARTIAL WIDTHS

$\Gamma(\pi\pi)$		$\Gamma_1$
VALUE (MeV)	EVTS	DOCUMENT ID
<b>157.2 ± 4.0 OUR FIT</b>		
<b>157.0 ± 6.0 −1.0</b>		
• • • We do not use the following data for averages, fits, limits, etc. • • •		
152 ± 8	870	<sup>2</sup> SCHEGELSKY 06A RVUE $\gamma\gamma \rightarrow K_S^0 \bar{K}_S^0$

$\Gamma(K\bar{K})$		$\Gamma_3$
VALUE (MeV)	EVTS	DOCUMENT ID
<b>8.5 ± 0.8 OUR FIT</b>		
<b>9.0 ± 0.7 −0.3</b>		
• • • We do not use the following data for averages, fits, limits, etc. • • •		
7.5 ± 2.0	870	<sup>2</sup> SCHEGELSKY 06A RVUE $\gamma\gamma \rightarrow K_S^0 \bar{K}_S^0$

$\Gamma(\eta\eta)$		$\Gamma_5$
VALUE (MeV)	EVTS	DOCUMENT ID
<b>0.75 ± 0.14 OUR FIT</b>		
<b>1.0 ± 0.1</b>		
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1.8 ± 0.4	870	<sup>2</sup> SCHEGELSKY 06A RVUE $\gamma\gamma \rightarrow K_S^0 \bar{K}_S^0$

$\Gamma(\gamma\gamma)$		$\Gamma_7$
The value of this width depends on the theoretical model used. Unitary approaches with scalars typically (with exception of PENNINGTON 08) give values clustering around 2.6 keV; without an S-wave contribution, values are systematically higher (typically around 3 keV).		

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.6 ± 0.5 OUR FIT</b>				Error includes scale factor of 1.4.
<b>2.93 ± 0.40</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.14 ± 0.20		<sup>3</sup> DAI 14A	RVUE	Compilation
3.82 ± 0.30		<sup>4,5</sup> PENNINGTON 08	RVUE	Compilation
2.55 ± 0.15	870	<sup>5,6</sup> PENNINGTON 08	RVUE	Compilation
2.84 ± 0.35		<sup>2</sup> SCHEGELSKY 06A	RVUE	$\gamma\gamma \rightarrow K_S^0 \bar{K}_S^0$
2.93 ± 0.23 ± 0.32		BOGLIONE 99	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-, \pi^0\pi^0$
2.58 ± 0.13 ± 0.36 −0.27		<sup>7</sup> YABUKI 95	VNS	
3.10 ± 0.35 ± 0.35		<sup>8</sup> BEHREND 92	CELL	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.27 ± 0.47 ± 0.11		<sup>9</sup> BLINOV 92	MD1	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
3.15 ± 0.04 ± 0.39		ADACHI 90D	TOPZ	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
3.19 ± 0.16 ± 0.29 −0.28		BOYER 90	MRK2	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.35 ± 0.65		MARSISKE 90	CBAL	$e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
3.19 ± 0.09 ± 0.22 −0.38	2177	<sup>10</sup> MORGAN 90	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-, \pi^0\pi^0$
3.2 ± 0.1 ± 0.4		JADE 90	RVUE	$e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
2.5 ± 0.1 ± 0.5		<sup>11</sup> AIHARA 86B	TPC	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.85 ± 0.25 ± 0.5		BEHREND 84B	CELL	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.70 ± 0.05 ± 0.20		<sup>12</sup> BERGER 84	PLUT	$e^+e^- \rightarrow e^+e^-2\pi$
2.52 ± 0.13 ± 0.38		COURAU 84	DLCO	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.7 ± 0.2 ± 0.6		<sup>13</sup> SMITH 84C	MRK2	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.9 ± 0.6 −0.4		EDWARDS 82F	CBAL	$e^+e^- \rightarrow e^+e^-2\pi^0$
3.2 ± 0.2 ± 0.6		<sup>14</sup> EDWARDS 82F	CBAL	$e^+e^- \rightarrow e^+e^-2\pi^0$
3.6 ± 0.3 ± 0.5		BRANDELIK 81B	TASS	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
2.3 ± 0.8		ROUSSARIE 81	MRK2	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
		<sup>15</sup> BERGER 80B	PLUT	$e^+e^-$

$\Gamma(e^+e^-)$					$\Gamma_{10}$	
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;0.11</b>	90	ACHASOV	00K	SND	$e^+e^- \rightarrow \pi^0\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<1.7	90	VOROBYEV	88	ND	$e^+e^- \rightarrow \pi^0\pi^0$	

- From a partial-wave analysis of data using a K-matrix formalism with 5 poles.
- From analysis of L3 data at 91 and 183–209 GeV and using SU(3) relations.
- Based on a K-matrix analysis of BELLE data from MORI 07, UEHARA 08A, UEHARA 09 and UEHARA 13. The width is derived for the pole on the third sheet which is closest to the physical axis. Supersedes PENNINGTON 08.
- Solution A (preferred solution based on  $\chi^2$ -analysis).
- Dispersion theory based amplitude analysis of BOYER 90, MARSISKE 90, BEHREND 92, and MORI 07.
- Solution B (worse than solution A; still acceptable when systematic uncertainties are included).
- With a narrow scalar state around 1220 MeV.
- Using a unitarized model with a 300–500 keV wide scalar at 1100 MeV.
- Using the unitarized model of LYTH 85.
- Error includes spread of different solutions. Data of MARK2 and CRYSTAL BALL used in the analysis. Authors report strong correlations with  $\gamma\gamma$  width of  $f_0(1370)$ :  $\Gamma(f_2) + 1/4 \Gamma(f^0) = 3.6 \pm 0.3$  KeV.
- Radiative corrections modify the partial widths; for instance the COURAU 84 value becomes  $2.66 \pm 0.21$  in the calculation of LANDRO 86.
- Using the MENNESSIER 83 model.
- Superseded by BOYER 90.
- If helicity = 2 assumption is not made.
- Using mass, width and  $B(f_2(1270) \rightarrow 2\pi)$  from PDG 78.

### $f_2(1270) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$			$\Gamma_3\Gamma_7/\Gamma$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>0.121±0.020 OUR FIT</b>	Error includes scale factor of 1.3.		
<b>0.091±0.007±0.027</b>	<sup>1</sup> ALBRECHT	90G ARG	$e^+e^- \rightarrow e^+e^-K^+K^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.104±0.007±0.072	<sup>2</sup> ALBRECHT	90G ARG	$e^+e^- \rightarrow e^+e^-K^+K^-$
<sup>1</sup> Using an incoherent background.			
<sup>2</sup> Using a coherent background.			

$\Gamma(\eta\eta) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_5\Gamma_7/\Gamma$		
VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b><math>11.5^{+1.8+4.5}_{-2.0-3.7}</math></b>	<sup>1</sup> UEHARA	10A BELL	$10.6\text{ e}^+\text{e}^- \rightarrow \text{e}^+\text{e}^-\eta\eta$
<sup>1</sup> including interference with the $f_2'(1525)$ (parameters fixed to the values from the 2008 edition of this review, PDG 08) and $f_0(Y)$ .			

See key on page 601

## Meson Particle Listings

 $f_2(1270)$ 

Helicity-0/Helicity-2 RATIO IN $\gamma\gamma \rightarrow f_2(1270) \rightarrow \pi\pi$				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
$3.7 \pm 0.3 \pm^{15.9}_{2.9}$	UEHARA	08A	BELL	$10.6 e^+e^- \rightarrow e^+e^- \pi^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$9.5 \pm 1.8$	<sup>1</sup> DAI	14A	RVUE	Compilation
13	<sup>2,3</sup> PENNINGTON 08	RVUE	Compilation	
26	<sup>3,4</sup> PENNINGTON 08	RVUE	Compilation	

<sup>1</sup> Based on a  $K$ -matrix analysis of BELLE data from MORI 07, UEHARA 08A, UEHARA 09 and UEHARA 13. The width is derived for the pole on the third sheet which is closest to the physical axis.

<sup>2</sup> Solution A (preferred solution based on  $\chi^2$ -analysis).

<sup>3</sup> Dispersion theory based amplitude analysis of BOYER 90, MARSISKE 90, BEHREND 92, and MORI 07.

<sup>4</sup> Solution B (worse than solution A; still acceptable when systematic uncertainties are included).

 $f_2(1270)$  BRANCHING RATIOS

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
$0.842 \pm 0.029$	<b>OUR FIT</b> Error includes scale factor of 1.1.				
$0.837 \pm 0.020$	<b>OUR AVERAGE</b>				
$0.849 \pm 0.025$		CHABAUD	83	ASPK	$17 \pi^- p$ polarized
$0.85 \pm 0.05$	250	BEAUPRE	71	HBC	$8 \pi^+ p \rightarrow \Delta^{++} f_2$
$0.8 \pm 0.04$	600	OH	70	HBC	$1.26 \pi^- p \rightarrow \pi^+ \pi^- n$

$\Gamma(\pi^+ \pi^- 2\pi^0)/\Gamma(\pi\pi)$					$\Gamma_2/\Gamma_1$
Should be twice $\Gamma(2\pi^+ 2\pi^-)/\Gamma(\pi\pi)$ if decay is $\rho\rho$ . (See ASCOLI 68d.)					
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	

$0.091 \pm 0.014$  **OUR FIT** Error includes scale factor of 1.2.

$0.15 \pm 0.06$  600 EISENBERG 74 HBC  $4.9 \pi^+ p \rightarrow \Delta^{++} f_2$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.07 EMMS 75D DBC  $4 \pi^+ n \rightarrow p f_2$

$\Gamma(K\bar{K})/\Gamma(\pi\pi)$					$\Gamma_3/\Gamma_1$
We average only experiments which either take into account $f_2(1270)$ - $a_2(1320)$ interference explicitly or demonstrate that $a_2(1320)$ production is negligible.					
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	

$0.054 \pm 0.005$  **OUR FIT** Error includes scale factor of 2.7.

$0.041 \pm 0.004$  **OUR AVERAGE**

$0.045 \pm 0.01$  <sup>1</sup>BARGIOTTI 03 OBLX  $\bar{p}p$

$0.037 \pm 0.008$  ET KIN 82b MPS  $23 \pi^- p \rightarrow n 2K_S^0$

$0.045 \pm 0.009$  CHABAUD 81 ASPK  $17 \pi^- p$  polarized

$0.039 \pm 0.008$  LOVERRE 80 HBC  $4 \pi^- p \rightarrow K\bar{K}N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.052 \pm 0.025$  ABLIKIM 04e BES2  $J/\psi \rightarrow \omega K^+ K^-$

$0.036 \pm 0.005$  <sup>2</sup>COSTA... 80 OMEG  $1-2.2 \pi^- p \rightarrow K^+ K^- n$

$0.030 \pm 0.005$  <sup>3</sup>MARTIN 79 RVUE

$0.027 \pm 0.009$  <sup>4</sup>POLYCHRO... 79 STRC  $7 \pi^- p \rightarrow n 2K_S^0$

$0.025 \pm 0.015$  EMMS 75D DBC  $4 \pi^+ n \rightarrow p f_2$

$0.031 \pm 0.012$  20 ADERHOLZ 69 HBC  $8 \pi^+ p \rightarrow K^+ K^- \pi^+ p$

$\Gamma(2\pi^+ 2\pi^-)/\Gamma(\pi\pi)$					$\Gamma_4/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
$0.033 \pm 0.005$	<b>OUR FIT</b> Error includes scale factor of 1.2.				
$0.033 \pm 0.004$	<b>OUR AVERAGE</b> Error includes scale factor of 1.1.				

$0.024 \pm 0.006$  160 EMMS 75D DBC  $4 \pi^+ n \rightarrow p f_2$

$0.051 \pm 0.025$  70 EISENBERG 74 HBC  $4.9 \pi^+ p \rightarrow \Delta^{++} f_2$

$0.043 \pm 0.007$  285 LOUIE 74 HBC  $3.9 \pi^- p \rightarrow n f_2$

$0.037 \pm 0.007$  154 ANDERSON 73 DBC  $6 \pi^+ n \rightarrow p f_2$

$0.047 \pm 0.013$  OH 70 HBC  $1.26 \pi^- p \rightarrow \pi^+ \pi^- n$

$\Gamma(\eta\eta)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>4.0 ± 0.8 OUR FIT</b>	Error includes scale factor of 2.1.			

$2.7 \pm 0.7$  BINON 05 GAMS  $33 \pi^- p \rightarrow \eta\eta n$

$2.8 \pm 0.7$  ALDE 86D GAM4  $100 \pi^- p \rightarrow 2\eta n$

$5.2 \pm 1.7$  BINON 83 GAM2  $38 \pi^- p \rightarrow 2\eta n$

$\Gamma(\eta\eta)/\Gamma(\pi\pi)$					$\Gamma_5/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$0.003 \pm 0.001$		BARBERIS	00E	$450 p\bar{p} \rightarrow p_f \eta p_S$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.05$  95 EDWARDS 82f CBAL  $e^+e^- \rightarrow e^+e^- 2\eta$

$<0.016$  95 EMMS 75D DBC  $4 \pi^+ n \rightarrow p f_2$

$<0.09$  95 EISENBERG 74 HBC  $4.9 \pi^+ p \rightarrow \Delta^{++} f_2$

$\Gamma(4\pi^0)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
$0.0030 \pm 0.0010$	<b>OUR FIT</b>				
$0.003 \pm 0.001$	$400 \pm 50$	ALDE	87	GAM4	$100 \pi^- p \rightarrow 4\pi^0 n$

$\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$1.57 \pm 0.01 \pm^{1.39}_{-0.14}$  UEHARA 08A BELL  $10.6 e^+e^- \rightarrow e^+e^- \pi^0 \pi^0$

$\Gamma(\eta\pi\pi)/\Gamma(\pi\pi)$					$\Gamma_8/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.010$	95	EMMS	75D	DBC	$4 \pi^+ n \rightarrow p f_2$

$\Gamma(K^0 K^- \pi^+ + \text{c.c.})/\Gamma(\pi\pi)$					$\Gamma_9/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.004$	95	EMMS	75D	DBC	$4 \pi^+ n \rightarrow p f_2$

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_{10}/\Gamma$
VALUE (units $10^{-10}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
$<6$	90	ACHASOV	00K	SND	$e^+e^- \rightarrow \pi^0 \pi^0$

<sup>1</sup> Coupled channel analysis of  $\pi^+ \pi^- \pi^0$ ,  $K^+ K^- \pi^0$ , and  $K^\pm K_S^0 \pi^\mp$ .

<sup>2</sup> Re-evaluated by CHABAUD 83.

<sup>3</sup> Includes PAWLICKI 77 data.

<sup>4</sup> Takes into account the  $f_2(1270)$ - $f_2'(1525)$  interference.

 $f_2(1270)$  REFERENCES

DOBBS	15	PR D91 052006	S. Dobbs <i>et al.</i>	(NWES)
DAI	14A	PR D90 036004	L.-Y. Dai, M.R. Pennington	(CEBAF)
BOGOLYUB...	13	PAN 76 1324	M.Yu. Bogolyubsky <i>et al.</i>	(HYPERON-M Collab.)
		Translated from YAF 76 1389.		
UEHARA	13	PTEP 2013 123C01	S. Uehara <i>et al.</i>	(BELLE Collab.)
UEHARA	10A	PR D82 114031	S. Uehara <i>et al.</i>	(BELLE Collab.)
ANISOVICH	09	IJMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
UEHARA	09	PR D79 052009	S. Uehara <i>et al.</i>	(BELLE Collab.)
PDG	08	PL B667 1	C. Amisler <i>et al.</i>	(PDG Collab.)
PENNINGTON	08	EPJ C56 1	M.R. Pennington <i>et al.</i>	
UEHARA	08A	PR D78 052004	S. Uehara <i>et al.</i>	(BELLE Collab.)
MORI	07	PR D75 051101	T. Mori <i>et al.</i>	(BELLE Collab.)
ABLIKIM	06V	PL B642 441	M. Ablikim <i>et al.</i>	(BES Collab.)
SCHEGELSKY	06A	EPJ A27 207	V.A. Schegelsky <i>et al.</i>	
ABLIKIM	05	PL B607 243	M. Ablikim <i>et al.</i>	(BES Collab.)
BINON	05	PAN 68 960	F. Binon <i>et al.</i>	
		Translated from YAF 68 998.		
ABLIKIM	04E	PL B603 138	M. Ablikim <i>et al.</i>	(BES Collab.)
BARGIOTTI	03	EPJ C26 371	M. Bargiotti <i>et al.</i>	(OBELIX Collab.)
TIKHOMIROV	03	PAN 66 828	G.D. Tikhomirov <i>et al.</i>	
		Translated from YAF 66 860.		
ACHASOV	00K	PL B492 8	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
BARBERIS	00E	PL B479 59	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BOGLIONE	99	EPJ C9 11	M. Boglione, M.R. Pennington	
ALDE	98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
		Also PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
		Translated from YAF 62 446.		
ALDE	97	PL B397 350	D.M. Alde <i>et al.</i>	(GAMS Collab.)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
GRYGOREV	96	PAN 59 2105	V.K. Grigoriev, O.N. Baloshin, B.P. Barkov	(ITEP)
		Translated from YAF 59 2187.		
YABUKI	95	JPSJ 64 435	F. Yabuki <i>et al.</i>	(VENUS Collab.)
PROKOSHNIK	94	SPD 39 420	Y.D. Prokoshkin, A.A. Kondashov	(SERP)
		Translated from DANS 336 613.		
BEHREND	92	ZPHY C56 381	H.J. Behrend	(CELLO Collab.)
BLINOV	92	ZPHY C53 33	A.E. Blinov <i>et al.</i>	(NOVO)
AGUILAR...	91	ZPHY C50 405	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
AKER	91	PL B260 249	E. Aker <i>et al.</i>	(Crystal Barrel Collab.)
ADACHI	90D	PL B234 185	I. Adachi <i>et al.</i>	(TOPAZ Collab.)
ALBRECHT	90G	ZPHY C48 183	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BOYER	90	PR D42 1350	J. Boyer <i>et al.</i>	(Mark II Collab.)
BREAKSTONE	90	ZPHY C48 569	A.M. Breakstone <i>et al.</i>	(ISU, BCGA, CERN+)
MARSISKE	90	PR D41 3324	H. Marsiske <i>et al.</i>	(Crystal Ball Collab.)
MORGAN	90	ZPHY C48 623	D. Morgan, M.R. Pennington	(RAL, DURH)
OEST	90	ZPHY C47 343	T. Oest <i>et al.</i>	(JADE Collab.)
AUGUSTIN	89	NP B320 1	J.E. Augustin, G. Cosme	(DM2 Collab.)
VOROBYEV	88	SJNP 48 273	P.V. Vorobyev <i>et al.</i>	(NOVO)
		Translated from YAF 48 436.		
ALDE	87	PL B198 286	D.M. Alde <i>et al.</i>	(LANL, BRUX, SERP, LAPP)
AUGUSTIN	87	ZPHY C36 369	J.E. Augustin <i>et al.</i>	(LALO, CLER, FRAS+)
ABACHI	86B	PRL 57 1990	S. Abachi <i>et al.</i>	(PURD, ANL, IND, MICH+)
AIHARA	86B	PRL 57 404	H. Aihara <i>et al.</i>	(TPC-2 $\gamma$ Collab.)
ALDE	86D	NP B269 485	D.M. Alde <i>et al.</i>	(BELG, LAPP, SERP, CERN+)
LANDRO	86	PL B172 445	M. Landro, K.J. Mork, H.A. Olsen	(UTRO)
LONGACRE	86	PL B177 223	R.S. Longacre <i>et al.</i>	(BNL, BRAN, CUNY+)
LYTH	85	JP G11 459	D.H. Lyth	
BEHREND	84B	ZPHY C23 223	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BERGER	84	ZPHY C26 199	C. Berger <i>et al.</i>	(PLUTO Collab.)
COURAU	84	PL 147B 227	A. Courau <i>et al.</i>	(CIT, SLAC)
SMITH	84C	PR D30 851	J.R. Smith <i>et al.</i>	(SLAC, LBL, HARV)
BINON	83	NC 78A 313	F.G. Binon <i>et al.</i>	(BELG, LAPP, SERP+)
		Also SJNP 38 561	F.G. Binon <i>et al.</i>	(BELG, LAPP, SERP+)
		Translated from YAF 38 934.		
CHABAUD	83	NP B223 1	V. Chabaud <i>et al.</i>	(CERN, CRAC, MPIM)
DENNEY	83	PR D28 2726	D.L. Denney <i>et al.</i>	(IOWA, MICH)
MENNESSIER	83	ZPHY C16 241	G. Mennessier	(MONP)
APEL	82	NP B201 197	W.D. Apel <i>et al.</i>	(KARLK, KARLE, PISA, SERP+)
CASON	82	PRL 48 1316	N.M. Cason <i>et al.</i>	(NDAM, ANL)
EDWARDS	82F	PL 110B 82	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
ETKIN	82B	PR D25 1786	A. Etkin <i>et al.</i>	(BNL, CUNY, TUFTS, VAND)
BRANDELIK	81B	ZPHY C10 117	R. Brandelik <i>et al.</i>	(TASSO Collab.)
CHABAUD	81	APP B12 575	V. Chabaud <i>et al.</i>	(CERN, CRAC, MPIM)
GIDAI	81	PL 107B 153	G. Gidai <i>et al.</i>	(SLAC, LBL)
ROUSSARIE	81	PL 105B 304	A. Roussarie <i>et al.</i>	(SLAC, LBL)
BERGER	80B	PL 94B 254	C. Berger <i>et al.</i>	(PLUTO Collab.)
COSTA...	80	NP B175 402	G. Costa de Beaumgard <i>et al.</i>	(BARI, BONN+)
LOVERRE	80	ZPHY C6 187	P.F. Loverre <i>et al.</i>	(CERN, CDEF, MADR+)
CORDEN	79	NP B157 250	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TELA+)
MARTIN	79	NP B158 520	A.D. Martin, E.N. Ozmutlu	(DURH)

Meson Particle Listings

$f_2(1270)$ ,  $f_1(1285)$

POLYCHRO...	79	PR D19 1317	V.A. Polychronakos <i>et al.</i>	(NDAM, ANL)
PDG	78	PL 75B 1	C. Bricman <i>et al.</i>	
ANTIPOV	77	NP B119 45	Y.M. Antipov <i>et al.</i>	(SERP, GEVA)
PAWLICKI	77	PR D15 3196	A.J. Pawlicki <i>et al.</i>	(ANL)
DEUTSCH...	76	NP B103 426	M. Deutschmann <i>et al.</i>	(AACH3, BERL, BONN+)
APEL	75	PL 57B 398	W.D. Apel <i>et al.</i>	(KARLK, KARLE, PISA, SERP+)
EMMS	75D	NP B96 155	M.J. Emms <i>et al.</i>	(BIRM, DURH, RHEL)
EISENBERG	74	PL 52B 239	Y. Eisenberg <i>et al.</i>	(REHO)
ENGLER	74	PL D10 2070	A. Engler <i>et al.</i>	(CMU, CASE)
LOUIE	74	PL 48B 385	J. Louie <i>et al.</i>	(SACL, CERN)
ANDERSON	73	PRL 31 562	J.C. Anderson <i>et al.</i>	(CMU, CASE)
TAKAHASHI	72	PR D6 1266	K. Takahashi <i>et al.</i>	(TOHOK, PENN, NDAM+)
BEAUPRE	71	NP B28 77	J.V. Beaupre <i>et al.</i>	(AACH, BERL, CERN)
FLATTE	71	PL 34B 551	S.M. Flatte <i>et al.</i>	(LBL)
ARMENISE	70	LNC 4 199	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ)
OH	70	PR D1 2494	B.Y. Oh <i>et al.</i>	(WISC, TNTO)JP
STUNTEBECK	70	PL 32B 391	P.H. Stuntebeck <i>et al.</i>	(NDAM)
ADERHOLZ	69	NP B11 259	M. Aderholz <i>et al.</i>	(AACH3, BERL, CERN+)
ARMENISE	68	NC 54A 999	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ+)
ASCOLI	68D	PRL 21 1712	G. Ascoli <i>et al.</i>	(ILL)
BOESEBECK	68	NP B4 501	K. Boesebeck <i>et al.</i>	(AACH, BERL, CERN)
JOHNSON	68	PR 176 1651	P.B. Johnson <i>et al.</i>	(NDAM, PURD, SLAC)
EISNER	67	PR 164 1699	R.L. Eisner <i>et al.</i>	(PURD)
DERADO	65	PRL 14 872	I. Derado <i>et al.</i>	(NDAM)
LEE	64	PRL 12 342	Y.Y. Lee <i>et al.</i>	(MICH)
BONDAR	63	PL 5 153	L. Bondar <i>et al.</i>	(AACH, BIRM, BONN, DESY+)

$f_1(1285)$

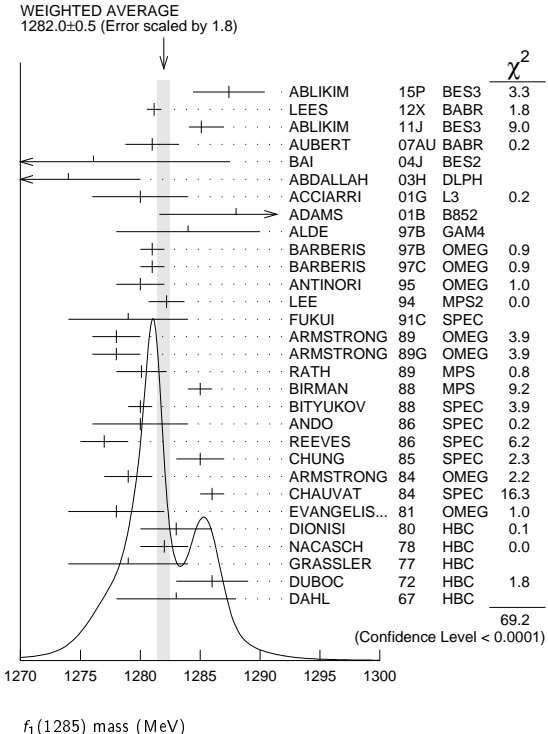
$I^G(J^{PC}) = 0^+(1^{++})$

$f_1(1285)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1282.0 ± 0.5</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.		
1287.4 ± 3.0	87	ABLIKIM	15P BES3	$J/\psi \rightarrow K^+ K^- 3\pi$
1281.16 ± 0.39 ± 0.45		<sup>1</sup> LEES	12X BABR	$\tau^- \rightarrow \pi^- f_1(1285) \nu_\tau$
1285.1 ± 1.0 ± 1.6 ± 0.3		<sup>2</sup> ABLIKIM	11J BES3	$J/\psi \rightarrow \omega(\eta\pi^+\pi^-)$
1281 ± 2 ± 1		AUBERT	07AU BABR	10.6 $e^+e^- \rightarrow f_1(1285)\pi^+\pi^-\gamma$
1276.1 ± 8.1 ± 8.0	203	BAI	04J BES2	$J/\psi \rightarrow \gamma\gamma\pi^+\pi^-$
1274 ± 6 ± 237		ABDALLAH	03H DLPH	91.2 $e^+e^- \rightarrow K_S^0 K^\pm \pi^\mp + X$
1280 ± 4		ACCIARRI	01G L3	
1288 ± 4 ± 5 ± 20k		ADAMS	01B B852	18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
1284 ± 6 ± 1400		ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta\pi^0\pi^0 n$
1281 ± 1		BARBERIS	97B OMEG	450 $pp \rightarrow pp2(\pi^+\pi^-)$
1281 ± 1		BARBERIS	97C OMEG	450 $pp \rightarrow pp K_S^0 K^\pm \pi^\mp$
1280 ± 2		<sup>3</sup> ANTINORI	95 OMEG	300,450 $pp \rightarrow pp2(\pi^+\pi^-)$
1282.2 ± 1.5		LEE	94 MPS2	18 $\pi^- p \rightarrow K^+ \bar{K}^0 2\pi^- p$
1279 ± 5		FUKUI	91C SPEC	8.95 $\pi^- p \rightarrow \eta\pi^+\pi^- n$
1278 ± 2	140	ARMSTRONG	89 OMEG	300 $pp \rightarrow K \bar{K} \pi p p$
1278 ± 2		ARMSTRONG	89G OMEG	85 $\pi^+ p \rightarrow 4\pi\pi p, pp \rightarrow 4\pi pp$
1280.1 ± 2.1	60	RATH	89 MPS	21.4 $\pi^- p \rightarrow K_S^0 K_S^0 \pi^0 n$
1285 ± 1	4750	<sup>4</sup> BIRMAN	88 MPS	8 $\pi^- p \rightarrow K^+ \bar{K}^0 \pi^- n$
1280 ± 1	504	BITYUKOV	88 SPEC	32.5 $\pi^- p \rightarrow K^+ K^- \pi^0 n$
1280 ± 4		ANDO	86 SPEC	8 $\pi^- p \rightarrow \eta\pi^+\pi^- n$
1277 ± 2	420	REEVES	86 SPEC	6.6 $p\bar{p} \rightarrow K K \pi X$
1285 ± 2		CHUNG	85 SPEC	8 $\pi^- p \rightarrow N K \bar{K} \pi$
1279 ± 2	604	ARMSTRONG	84 OMEG	85 $\pi^+ p \rightarrow K \bar{K} \pi p p, pp \rightarrow K \bar{K} \pi p p$
1286 ± 1		CHAUVAT	84 SPEC	ISR 31.5 $pp$
1278 ± 4		EVANGELIS...	81 OMEG	12 $\pi^- p \rightarrow \eta\pi^+\pi^-\pi^- p$
1283 ± 3	103	DIONISI	80 HBC	4 $\pi^- p \rightarrow K \bar{K} \pi n$
1282 ± 2	320	NACASCH	78 HBC	0.7,0.76 $\bar{p} p \rightarrow K \bar{K} 3\pi$
1279 ± 5	210	GRASSLER	77 HBC	16 $\pi^- p$
1286 ± 3	180	DUBOC	72 HBC	1.2 $\bar{p} p \rightarrow 2K4\pi$
1283 ± 5		DAHL	67 HBC	1.6-4.2 $\pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1284.2 ± 2.2		<sup>5</sup> AAIJ	14Y LHCB	$\bar{B}^0(s) \rightarrow J/\psi 2(\pi^+\pi^-)$
1281.9 ± 0.5		<sup>5</sup> SOSA	99 SPEC	$pp \rightarrow p_{\text{slow}} (K_S^0 K^\pm \pi^\mp) p_{\text{fast}}$
1282.8 ± 0.6		<sup>5</sup> SOSA	99 SPEC	$pp \rightarrow p_{\text{slow}} (K_S^0 K^- \pi^+) p_{\text{fast}}$
1270 ± 10		AMELIN	95 VES	37 $\pi^- N \rightarrow \pi^- \pi^+ \pi^- \gamma N$
1280 ± 2		ABATZIS	94 OMEG	450 $pp \rightarrow pp2(\pi^+\pi^-)$
1282 ± 4		ARMSTRONG	93C E760	$\bar{p} p \rightarrow \pi^0 \eta \eta \rightarrow 6\gamma$
1270 ± 6 ± 10		ARMSTRONG	92C OMEG	300 $pp \rightarrow pp\pi^+\pi^-\gamma$
1281 ± 1		ARMSTRONG	89E OMEG	300 $pp \rightarrow pp2(\pi^+\pi^-)$
1279 ± 6 ± 10 ± 16		BECKER	87 MRK3	$e^+e^- \rightarrow \phi K \bar{K} \pi$
1286 ± 9		GIDAL	87 MRK2	$e^+e^- \rightarrow e^+e^-\eta\pi^+\pi^-$
1287 ± 5 ± 1279	353	<sup>6</sup> BITYUKOV	84B SPEC	32 $\pi^- p \rightarrow K^+ K^- \pi^0 n$
		TORNQVIST	82B RVUE	

1275 ± 6	31	BROMBERG	80 SPEC	100 $\pi^- p \rightarrow K \bar{K} \pi X$
1288 ± 9	200	GURTU	79 HBC	4.2 $K^- p \rightarrow n\eta 2\pi$
~ 1275.0	46	<sup>7</sup> STANTON	79 CNTR	8.5 $\pi^- p \rightarrow n2\gamma 2\pi$
1271 ± 10	34	CORDEN	78 OMEG	12-15 $\pi^- p \rightarrow K^+ K^- \pi n$
1295 ± 12	85	CORDEN	78 OMEG	12-15 $\pi^- p \rightarrow n5\pi$
1292 ± 10	150	DEFOIX	72 HBC	0.7 $\bar{p} p \rightarrow 7\pi$
1280 ± 3	500	<sup>8</sup> THUN	72 MMS	13.4 $\pi^- p$
1303 ± 8		BARDADIN...	71 HBC	8 $\pi^+ p \rightarrow p6\pi$
1283 ± 6		BOESEBECK	71 HBC	16.0 $\pi p \rightarrow p5\pi$
1270 ± 10		CAMPBELL	69 DBC	2.7 $\pi^+ d$
1285 ± 7		LORSTAD	69 HBC	0.7 $\bar{p} p$ , 4,5-body
1290 ± 7		D'ANDLAU	68 HBC	1.2 $\bar{p} p$ , 5-6 body

- <sup>1</sup> Using the  $2\pi+2\pi^-$  and  $\pi^+\pi^-\eta$  modes of  $f_1(1285)$  decay.  
<sup>2</sup> The selected process is  $J/\psi \rightarrow \omega a_0(980)\pi$ .  
<sup>3</sup> Supersedes ABATZIS 94, ARMSTRONG 89E.  
<sup>4</sup> From partial wave analysis of  $K^+ \bar{K}^0 \pi^-$  system.  
<sup>5</sup> No systematic error given.  
<sup>6</sup> From a unitarized quark-model calculation.  
<sup>7</sup> From phase shift analysis of  $\eta\pi^+\pi^-$  system.  
<sup>8</sup> Seen in the missing mass spectrum.



$f_1(1285)$  WIDTH

Only experiments giving width error less than 20 MeV are kept for averaging.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>24.1 ± 1.0</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.		
18.3 ± 6.3	87	ABLIKIM	15P BES3	$J/\psi \rightarrow K^+ K^- 3\pi$
22.0 ± 3.1 ± 2.0 ± 1.5		<sup>1</sup> ABLIKIM	11J BES3	$J/\psi \rightarrow \omega(\eta\pi^+\pi^-)$
35 ± 6 ± 4		AUBERT	07AU BABR	10.6 $e^+e^- \rightarrow f_1(1285)\pi^+\pi^-\gamma$
40.0 ± 8.6 ± 9.3	203	BAI	04J BES2	$J/\psi \rightarrow \gamma\gamma\pi^+\pi^-$
29 ± 12 ± 237		ABDALLAH	03H DLPH	91.2 $e^+e^- \rightarrow K_S^0 K^\pm \pi^\mp + X$
45 ± 9 ± 7 ± 20k		ADAMS	01B B852	18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
55 ± 18 ± 24 ± 3	1400	ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta\pi^0\pi^0 n$
20 ± 2		BARBERIS	97B OMEG	450 $pp \rightarrow pp2(\pi^+\pi^-)$
36 ± 5		BARBERIS	97C OMEG	450 $pp \rightarrow pp K_S^0 K^\pm \pi^\mp$
		<sup>2</sup> ANTINORI	95 OMEG	300,450 $pp \rightarrow pp2(\pi^+\pi^-)$
29.0 ± 4.1		LEE	94 MPS2	18 $\pi^- p \rightarrow K^+ \bar{K}^0 2\pi^- p$
25 ± 4 ± 140		ARMSTRONG	89 OMEG	300 $pp \rightarrow K \bar{K} \pi p p$
22 ± 2 ± 4750		<sup>3</sup> BIRMAN	88 MPS	8 $\pi^- p \rightarrow K^+ \bar{K}^0 \pi^- n$

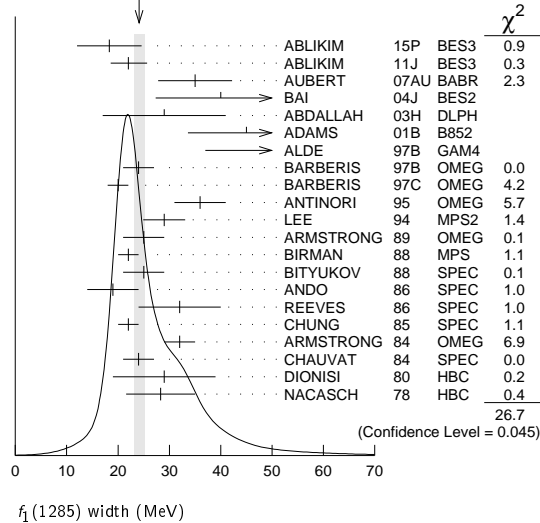


See key on page 601

## Meson Particle Listings

 $f_1(1285)$ 

25 ± 4	504	BITYUKOV	88	SPEC	32.5 $\pi^- p \rightarrow K^+ K^- \pi^0 n$
19 ± 5		ANDO	86	SPEC	8 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
32 ± 8	420	REEVES	86	SPEC	6.6 $p \bar{p} \rightarrow K K \pi X$
22 ± 2		CHUNG	85	SPEC	8 $\pi^- p \rightarrow N K \bar{K} \pi$
32 ± 3	604	ARMSTRONG	84	OMEG	85 $\pi^+ p \rightarrow K \bar{K} \pi \pi p$ , $p p \rightarrow K \bar{K} \pi p p$
24 ± 3		CHAUVAT	84	SPEC	ISR 31.5 $p p$
29 ± 10	103	DIONISI	80	HBC	4 $\pi^- p \rightarrow K \bar{K} \pi n$
28.3 ± 6.7	320	NACASCH	78	HBC	0.7, 0.76 $\bar{p} p \rightarrow K \bar{K} 3\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
32.4 ± 5.8	4	AAIJ	14	LHCb	$\bar{B}_{(s)}^0 \rightarrow J/\psi 2(\pi^+ \pi^-)$
18.2 ± 1.2	4	SOSA	99	SPEC	$p p \rightarrow p_{\text{slow}} (K_S^0 K^+ \pi^-)$
19.4 ± 1.5	4	SOSA	99	SPEC	$p p \rightarrow p_{\text{slow}} (K_S^0 K^- \pi^+)$
40 ± 5		ABATZIS	94	OMEG	450 $p p \rightarrow p p 2(\pi^+ \pi^-)$
31 ± 5		ARMSTRONG	89E	OMEG	300 $p p \rightarrow p p 2(\pi^+ \pi^-)$
41 ± 12		ARMSTRONG	89G	OMEG	85 $\pi^+ p \rightarrow 4\pi \pi p$ , $p p \rightarrow 4\pi p p$
17.9 ± 10.9	60	RATH	89	MPS	21.4 $\pi^- p \rightarrow K_S^0 K_S^0 \pi^0 n$
14 $^{+20}_{-14}$ ± 10	16	BECKER	87	MRK3	$e^+ e^- \rightarrow \phi K \bar{K} \pi$
26 ± 12		EVANGELIS...	81	OMEG	12 $\pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$
25 ± 15	200	GURTU	79	HBC	4.2 $K^- p \rightarrow n \eta 2\pi$
~ 10	5	STANTON	79	CNTR	8.5 $\pi^- p \rightarrow n 2\gamma 2\pi$
24 ± 18	210	GRASSLER	77	HBC	16 $\pi^\mp p$
28 ± 5	150	DEFOIX	72	HBC	0.7 $\bar{p} p \rightarrow 7\pi$
46 ± 9	180	DUBOC	72	HBC	1.2 $\bar{p} p \rightarrow 2K 4\pi$
37 ± 5	500	THUN	72	MMS	13.4 $\pi^- p$
10 ± 10		BOESEBECK	71	HBC	16.0 $p p \rightarrow p 5\pi$
30 ± 15		CAMPBELL	69	DBC	2.7 $\pi^+ d$
60 ± 15		LORSTAD	69	HBC	0.7 $\bar{p} p$ , 4,5-body
35 ± 10		DAHL	67	HBC	1.6–4.2 $\pi^- p$

<sup>1</sup> The selected process is  $J/\psi \rightarrow \omega a_0(980)\pi$ .<sup>2</sup> Supersedes ABATZIS 94, ARMSTRONG 89E.<sup>3</sup> From partial wave analysis of  $K^+ \bar{K}^0 \pi^-$  system.<sup>4</sup> No systematic error given.<sup>5</sup> From phase shift analysis of  $\eta \pi^+ \pi^-$  system.<sup>6</sup> Resolution is not unfolded.<sup>7</sup> Seen in the missing mass spectrum.WEIGHTED AVERAGE  
24.1 ± 1.0 (Error scaled by 1.3) $f_1(1285)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $4\pi$	$(33.1^{+2.1}_{-1.8})\%$	S=1.3
$\Gamma_2$ $\pi^0 \pi^0 \pi^+ \pi^-$	$(22.0^{+1.4}_{-1.2})\%$	S=1.3
$\Gamma_3$ $2\pi^+ 2\pi^-$	$(11.0^{+0.7}_{-0.6})\%$	S=1.3
$\Gamma_4$ $\rho^0 \pi^+ \pi^-$	$(11.0^{+0.7}_{-0.6})\%$	S=1.3
$\Gamma_5$ $\rho^0 \rho^0$	seen	
$\Gamma_6$ $4\pi^0$	$< 7 \times 10^{-4}$	CL=90%
$\Gamma_7$ $\eta \pi^+ \pi^-$	$(35 \pm 15)\%$	
$\Gamma_8$ $\eta \pi \pi$	$(52.4^{+1.9}_{-2.2})\%$	S=1.2

$\Gamma_9$	$a_0(980)\pi$ [ignoring $a_0(980) \rightarrow K \bar{K}$ ]	$(36 \pm 7)\%$	
$\Gamma_{10}$	$\eta \pi \pi$ [excluding $a_0(980)\pi$ ]	$(16 \pm 7)\%$	
$\Gamma_{11}$	$K \bar{K} \pi$	$(9.0 \pm 0.4)\%$	S=1.1
$\Gamma_{12}$	$K \bar{K}^*(892)$	not seen	
$\Gamma_{13}$	$\pi^+ \pi^- \pi^0$	$(3.0 \pm 0.9) \times 10^{-3}$	
$\Gamma_{14}$	$\rho^\pm \pi^\mp$	$< 3.1 \times 10^{-3}$	CL=95%
$\Gamma_{15}$	$\gamma \rho^0$	$(5.5 \pm 1.3)\%$	S=2.8
$\Gamma_{16}$	$\phi \gamma$	$(7.4 \pm 2.6) \times 10^{-4}$	
$\Gamma_{17}$	$\gamma \gamma^*$		
$\Gamma_{18}$	$\gamma \gamma$		

## CONSTRAINED FIT INFORMATION

An overall fit to 7 branching ratios uses 16 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 24.7$  for 12 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_9$	-17			
$x_{10}$	-8	-95		
$x_{11}$	46	-9	-4	
$x_{15}$	-36	-4	-2	-34
	$x_1$	$x_9$	$x_{10}$	$x_{11}$

 $f_1(1285) \Gamma(i) \Gamma(\gamma \gamma) / \Gamma_{\text{total}}$ 

VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.62</b>	95	GIDAL	87	MRK2 $e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.4 ± 0.4 OUR AVERAGE</b>				Error includes scale factor of 1.4.

1.18 ± 0.25 ± 0.20	26	1,2 AIHARA	88b	TPC $e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$
2.30 ± 0.61 ± 0.42		1,3 GIDAL	87	MRK2 $e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.8 ± 0.3 ± 0.3	420	4 ACHARD	02b L3	183–209 $e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$
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<sup>1</sup> Assuming a  $\rho$ -pole form factor.<sup>2</sup> Published value multiplied by  $\eta \pi \pi$  branching ratio 0.49.<sup>3</sup> Published value divided by 2 and multiplied by the  $\eta \pi \pi$  branching ratio 0.49.<sup>4</sup> Published value multiplied by the  $\eta \pi \pi$  branching ratio 0.52. $f_1(1285)$  BRANCHING RATIOS

$\Gamma(K \bar{K} \pi) / \Gamma(4\pi)$	$\Gamma_{11} / \Gamma_1$
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VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.271 ± 0.016 OUR FIT</b>			Error includes scale factor of 1.3.
<b>0.271 ± 0.016 OUR AVERAGE</b>			Error includes scale factor of 1.2.
0.265 ± 0.014	1 BARBERIS	97c OMEG	450 $p p \rightarrow p p K_S^0 K^\pm \pi^\mp$
0.28 ± 0.05	2 ARMSTRONG	89E OMEG	300 $p p \rightarrow p p f_1(1285)$
0.37 ± 0.03 ± 0.05	3 ARMSTRONG	89G OMEG	85 $\pi p \rightarrow 4\pi X$

<sup>1</sup> Using  $2(\pi^+ \pi^-)$  data from BARBERIS 97b.<sup>2</sup> Assuming  $\rho \pi \pi$  and  $a_0(980)\pi$  intermediate states.<sup>3</sup>  $4\pi$  consistent with being entirely  $\rho \pi \pi$ .

$\Gamma(\pi^0 \pi^0 \pi^+ \pi^-) / \Gamma_{\text{total}}$	$\Gamma_2 / \Gamma = \frac{2}{3} \Gamma_1 / \Gamma$
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VALUE	DOCUMENT ID
<b>0.220 <math>^{+0.014}_{-0.012}</math> OUR FIT</b>	

$\Gamma(2\pi^+ 2\pi^-) / \Gamma_{\text{total}}$	$\Gamma_3 / \Gamma = \frac{1}{3} \Gamma_1 / \Gamma$
---	---

VALUE	DOCUMENT ID
<b>0.110 <math>^{+0.007}_{-0.006}</math> OUR FIT</b>	

$\Gamma(\rho^0 \pi^+ \pi^-) / \Gamma_{\text{total}}$	$\Gamma_4 / \Gamma = \frac{1}{3} \Gamma_1 / \Gamma$
--	---

VALUE	DOCUMENT ID
<b>0.110 <math>^{+0.007}_{-0.006}</math> OUR FIT</b>	

$\Gamma(\rho^0 \pi^+ \pi^-) / \Gamma(2\pi^+ 2\pi^-)$	$\Gamma_4 / \Gamma_3$
--	-----------------------

VALUE	DOCUMENT ID	TECN	COMMENT
1.0 ± 0.4	GRASSLER	77 HBC	16 GeV $\pi^\pm p$

# Meson Particle Listings

## $f_1(1285)$

$\Gamma(\rho^0\rho^0)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	BARBERIS	00c	450 $p p \rightarrow p_f 4\pi p_S$	

$\Gamma(4\pi^0)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<7	ALDE	87	GAM4 100 $\pi^- p \rightarrow 4\pi^0 n$	

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\eta\pi^+\pi^-)$				$\Gamma_{13}/\Gamma_7$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
$0.86\pm 0.16\pm 0.20$	1 DOROFEEV	11	VES $\pi^- N \rightarrow \pi^- f_1(1285) N$	

<sup>1</sup> Value obtained selecting the region corresponding to  $f_0(980)$  in the  $\pi^+\pi^-$  mass spectrum.

$\Gamma(\eta\pi\pi)/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma = (\Gamma_9+\Gamma_{10})/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.524^{+0.019}_{-0.022}$ OUR FIT				Error includes scale factor of 1.2.

$\Gamma(4\pi)/\Gamma(\eta\pi\pi)$				$\Gamma_1/\Gamma_8 = \Gamma_1/(\Gamma_9+\Gamma_{10})$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.63\pm 0.06$ OUR FIT				Error includes scale factor of 1.2.
$0.41\pm 0.14$ OUR AVERAGE				

0.37±0.11±0.11	BOLTON	92	MRK3 $J/\psi \rightarrow \gamma f_1(1285)$	
0.64±0.40	GURTU	79	HBC 4.2 $K^- p$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.93±0.30	1 GRASSLER	77	HBC 16 $\pi^\mp p$	

<sup>1</sup> Assuming  $\rho\pi\pi$  and  $a_0(980)\pi$  intermediate states.

$\Gamma(2\pi^+2\pi^-)/\Gamma(\eta\pi\pi)$				$\Gamma_3/\Gamma_8$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.28\pm 0.02\pm 0.02$	1 LEES	12x	BABR $\tau^- \rightarrow \pi^- f_1(1285) \nu_\tau$	

<sup>1</sup> Assuming  $B(f_1(1285) \rightarrow \pi\pi\eta) = 3/2 B(f_1(1285) \rightarrow \pi^+\pi^-\eta)$ .

$\Gamma(a_0(980)\pi \text{ [ignoring } a_0(980) \rightarrow K\bar{K}])/ \Gamma(\eta\pi\pi)$				$\Gamma_9/\Gamma_8 = \Gamma_9/(\Gamma_9+\Gamma_{10})$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.69\pm 0.13$ OUR FIT				

$0.69^{+0.13}_{-0.12}$  OUR AVERAGE

0.72±0.15	GURTU	79	HBC 4.2 $K^- p$	
0.6 $^{+0.3}_{-0.2}$	CORDEN	78	OMEG 12-15 $\pi^- p$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>0.69	95	318	ACHARD 02b L3 183-209 $e^+e^- \rightarrow e^+e^-\eta\pi^+\pi^-$	
0.28±0.07	1400	ALDE	97b GAM4 100 $\pi^- p \rightarrow \eta\pi^0\pi^0 n$	
1.0 ±0.3	GRASSLER	77	HBC 16 $\pi^\mp p$	

$\Gamma(K\bar{K}\pi)/\Gamma(\eta\pi\pi)$				$\Gamma_{11}/\Gamma_8 = \Gamma_{11}/(\Gamma_9+\Gamma_{10})$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.171\pm 0.013$ OUR FIT				Error includes scale factor of 1.1.
$0.170\pm 0.012$ OUR AVERAGE				

0.166±0.01 ±0.008	BARBERIS	98c	OMEG 450 $p p \rightarrow p_f f_1(1285) p_S$	
0.42 ±0.15	GURTU	79	HBC 4.2 $K^- p$	
0.5 ±0.2	1 CORDEN	78	OMEG 12-15 $\pi^- p$	
0.20 ±0.08	2 DEFOIX	72	HBC 0.7 $\bar{p} p \rightarrow 7\pi$	
0.16 ±0.08	CAMPBELL	69	DBC 2.7 $\pi^+ d$	

<sup>1</sup> CORDEN 78 assumes low-mass  $\eta\pi\pi$  region is dominantly  $1^{++}$ . See BARBERIS 98c and MANAK 00a for discussion.

<sup>2</sup>  $K\bar{K}$  system characterized by the  $l = 1$  threshold enhancement. (See under  $a_0(980)$ ).

$\Gamma(K\bar{K}^*(892))/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
not seen	NACASCH	78	HBC 0.7,0.76 $\bar{p} p \rightarrow K\bar{K}^* 3\pi$	

• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	1 ACHARD	07	L3 183-209 $e^+e^- \rightarrow e^+e^-K_S^0 K^\pm \pi^\mp$	

<sup>1</sup> A clear signal of  $19.8 \pm 4.4$  events observed at high  $Q^2$ .

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
$0.30\pm 0.055\pm 0.074$	1 DOROFEEV	11	VES $\pi^- N \rightarrow \pi^- f_1(1285) N$	

<sup>1</sup> Value obtained selecting the region corresponding to  $f_0(980)$  in the  $\pi^+\pi^-$  mass spectrum. The systematic error includes the uncertainty on the partial width  $f_1 \rightarrow \eta\pi\pi$  obtained from PDG 10 data.

$\Gamma(\rho^\pm\pi^\mp)/\Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<0.31	95	DOROFEEV 11	VES	$\pi^- N \rightarrow \pi^- f_1(1285) N$

$\Gamma(\gamma\rho^0)/\Gamma_{\text{total}}$				$\Gamma_{15}/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
$5.5\pm 1.3$ OUR FIT				Error includes scale factor of 2.8.
$2.8\pm 0.7\pm 0.6$	AMELIN	95	VES 37 $\pi^- N \rightarrow \pi^- \pi^+ \pi^- \gamma N$	

• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5	95	BITYUKOV	91B SPEC	32 $\pi^- p \rightarrow \pi^+ \pi^- \gamma n$

$\Gamma(\gamma\rho^0)/\Gamma(2\pi^+2\pi^-)$				$\Gamma_{15}/\Gamma_3 = \Gamma_{15}/\frac{1}{3}\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.50\pm 0.13$ OUR FIT				Error includes scale factor of 2.5.
$0.45\pm 0.18$	1 COFFMAN	90	MRK3 $J/\psi \rightarrow \gamma\gamma\pi^+\pi^-$	

<sup>1</sup> Using  $B(J/\psi \rightarrow \gamma f_1(1285) \rightarrow \gamma\gamma\rho^0) = 0.25 \times 10^{-4}$  and  $B(J/\psi \rightarrow \gamma f_1(1285) \rightarrow \gamma 2\pi^+ 2\pi^-) = 0.55 \times 10^{-4}$  given by MIR 88.

$\Gamma(\eta\pi\pi)/\Gamma(\gamma\rho^0)$				$\Gamma_8/\Gamma_{15} = (\Gamma_9+\Gamma_{10})/\Gamma_{15}$
VALUE	DOCUMENT ID	TECN	COMMENT	
$9.5\pm 2.0$ OUR FIT				Error includes scale factor of 2.5.
$7.9\pm 0.9$ OUR AVERAGE				

10.0±1.0±2.0	BARBERIS	98c	OMEG 450 $p p \rightarrow p_f f_1(1285) p_S$	
7.5±1.0	1 ARMSTRONG	92c	OMEG 300 $p p \rightarrow p p \pi^+ \pi^- \gamma, p p \eta \pi^+ \pi^-$	

<sup>1</sup> Published value multiplied by 1.5.

$\Gamma(\gamma\rho^0)/\Gamma(K\bar{K}\pi)$					$\Gamma_{15}/\Gamma_{11}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					

<sup>1</sup> Using  $B(J/\psi \rightarrow \gamma f_1(1285) \rightarrow \gamma\gamma\rho^0) = 0.25 \times 10^{-4}$  and  $B(J/\psi \rightarrow \gamma f_1(1285) \rightarrow \gamma K\bar{K}\pi) < 0.72 \times 10^{-3}$ .

$\Gamma(\phi\gamma)/\Gamma(K\bar{K}\pi)$					$\Gamma_{16}/\Gamma_{11}$
VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$0.82\pm 0.21\pm 0.20$		19	BITYUKOV	88	SPEC 32.5 $\pi^- p \rightarrow$

• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.50	95	BARBERIS	98c	OMEG 450 $p p \rightarrow p_f f_1(1285) p_S$
<0.93	95	AMELIN	95	VES 37 $\pi^- N \rightarrow \pi^- \pi^+ \pi^- \gamma N$

## $f_1(1285)$ REFERENCES

ABLIKIM	15P	PR D92 012007	M. Ablikim <i>et al.</i>	(BES III Collab.)
AJJ	14Y	PRL 112 091802	R. Aaij <i>et al.</i>	(LHCb Collab.)
LEES	12X	PR D86 092010	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABLIKIM	11J	PRL 107 182001	M. Ablikim <i>et al.</i>	(BES III Collab.)
DOROFEEV	11	EPJ A47 68	V. Dorofeev <i>et al.</i>	(SERP, MIPT)
PDG	10	JP G37 075021	K. Nakamura <i>et al.</i>	(PDG Collab.)
ACHARD	07	JHEP 0703 018	P. Achard <i>et al.</i>	(L3 Collab.)
AUBERT	07AU	PR D76 092005	B. Aubert <i>et al.</i>	(BABAR Collab.)
BAI	04J	PL B594 47	J.Z. Bai <i>et al.</i>	(BES Collab.)
ABDALLAH	03H	PL B569 129	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ACHARD	02B	PL B526 269	P. Achard <i>et al.</i>	(L3 Collab.)
ACCIARRI	01B	PL B501 1	M. Acciarri <i>et al.</i>	(L3 Collab.)
ADAMS	01B	PL B516 264	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)
BARBERIS	00C	PL B471 440	D. Barberis <i>et al.</i>	(WA 102 Collab.)
MANAK	00A	PR D62 012003	J.J. Manak <i>et al.</i>	(BNL E852 Collab.)
SOSA	99	PRL 83 913	M. Sosa <i>et al.</i>	
BARBERIS	98C	PL B440 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ALDE	97B	PAN 60 386	D. Alde <i>et al.</i>	(GAMS Collab.)
BARBERIS	97B	Translated from YAF 60 458	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	97C	PL B413 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
AMELIN	95	ZPHY C66 71	D.V. Amelin <i>et al.</i>	(VES Collab.)
ANTINORI	95	PL B353 589	F. Antinori <i>et al.</i>	(ATHU, BARI, BIRM+)
ABATZIS	94	PL B324 509	S. Abatzis <i>et al.</i>	(ATHU, BARI, BIRM+)
LEE	94	PL B323 227	J.H. Lee <i>et al.</i>	(BNL, IND, KYUN, MASD+)
ARMSTRONG	93C	PL B307 394	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
ARMSTRONG	92C	ZPHY C54 371	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
BOLTON	92	PL B278 495	T. Bolton <i>et al.</i>	(Mark III Collab.)
BITYUKOV	91B	SJNP 54 318	S.I. Bityukov <i>et al.</i>	(SERP)
FUKUI	91C	PL B267 293	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
COFFMAN	90	PR D41 1410	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
ARMSTRONG	89	PL B221 216	T.A. Armstrong <i>et al.</i>	(CERN, CDEF, BIRM+)
ARMSTRONG	89E	PL B228 536	T.A. Armstrong, M. Benayoun	(ATHU, BARI, BIRM+)
ARMSTRONG	89G	ZPHY C43 55	T.A. Armstrong <i>et al.</i>	(CERN, BIRM, BARI+)
RATH	89	PR D40 693	M.G. Rath <i>et al.</i>	(NDAM, BRAN, BNL, CUNY+)
AIHARA	88B	PL B209 107	H. Aihara <i>et al.</i>	(TPC-2γ Collab.)
BIRMAN	88	PRL 61 1557	A. Birman <i>et al.</i>	(BNL, FSU, IND, MASD)
BITYUKOV	88	PL B203 327	S.I. Bityukov <i>et al.</i>	(SERP)
MIR	88	Photon-Photon 88, 126	R. Mir	(Mark III Collab.)
ALDE	87	PL B198 286	D.M. Alde <i>et al.</i>	(LANL, BRUX, SERP, LAPP)
BECKER	87	PRL 59 186	J.J. Becker <i>et al.</i>	(Mark III Collab.)
GIDAL	87	PRL 59 2012	G. Gidal <i>et al.</i>	(LBL, SLAC, HARV)
ANDO	86	PRL 57 1296	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
REEVES	86	PR D34 1960	D.F. Reeves <i>et al.</i>	(FLOR, BNL, IND+)
CHUNG	85	PRL 55 779	S.U. Chung <i>et al.</i>	(BNL, FLOR, IND+)
ARMSTRONG	84B	PL 146B 273	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
BITYUKOV	84B	PL 144B 133	S.I. Bityukov <i>et al.</i>	(SERP)
CHAUVAET	84	PL 145B 302	P. Chauvaet <i>et al.</i>	(CERN, CLER, UCLA+)
TORNQVIST	82B	NP B203 268	N.A. Tornqvist	(HELS)
EVANGELISTA	81	NP B178 197	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
BROMBERG	80	PR D22 1513	C.M. Bromberg <i>et al.</i>	(CIT, FNAL, ILLC+)
DIONISI	80	NP B169 1	C. Dionisi <i>et al.</i>	(CERN, MADR, CDEF+)
GURTU	79	NP B151 181	A. Gurtu <i>et al.</i>	(CERN, ZEEM, NIJM, OXF)
STANTON	79	PRL 42 346	N.R. Stanton <i>et al.</i>	(OSU, CARL, MCGI+)
CORDEN	78	NP B144 253	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TELA+)
NACASCH	78	NP B135 203	R. Nacasch <i>et al.</i>	(PARIS, MADR, CERN)
GRASSLER	77	NP B121 189	H. Grassler <i>et al.</i>	(AACH3, BERL, BONN+)
DEFOIX	72	NP B44 125	C. Defoix <i>et al.</i>	(CDEF, CERN)
DUBOC	72	NP B46 429	J. Duboc <i>et al.</i>	(PARIS, LVP)
THUN	72	PRL 28 1733	R. Thun <i>et al.</i>	(STON, NEAS)
BARDADIN-...	71	PR D4 2711	M. Bardadin-Otwinowska <i>et al.</i>	(WARS)
BOESEBECK	71	PL 34B 659	K. Boesebeck	(AACH, BERL, BONN, CERN, CRAC+)
CAMPBELL	69	PRL 22 1204	J.H. Campbell <i>et al.</i>	(PURD)
LORSTAD	69	NP B14 63	B. Lorstad <i>et al.</i>	(CDEF, CERN)
D'ANDLAU	68	NP B5 693	C. D'Andlau <i>et al.</i>	(CDEF, CERN, IRAD+)
DAHL	67	PR 163 1377	O.I. Dahl <i>et al.</i>	(LRL)

See key on page 601

## Meson Particle Listings

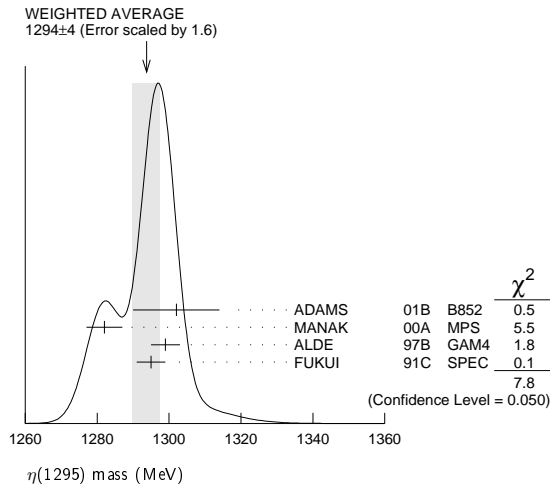
 $\eta(1295)$ ,  $\pi(1300)$  $\eta(1295)$ 

$$J^G(J^{PC}) = 0^+(0^{-+})$$

See also the mini-review under  $\eta(1405)$  $\eta(1295)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1294±4 OUR AVERAGE</b>	Error includes scale factor of 1.6. See the ideogram below.			
1302±9±8	20k	ADAMS	01B B852	18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
1282±5	9082	MANAK	00A MPS	18 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
1299±4	2100	ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta \pi^0 \pi^0 n$
1295±4		FUKUI	91C SPEC	8.95 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- |        |   |          |         |  |
|--------|---|----------|---------|--|
| 1264±8 | 1 | AUGUSTIN | 90 DM2  | $J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$ |
| ~ 1275 |   | STANTON  | 79 CNTR | 8.4 $\pi^- p \rightarrow n \eta 2\pi$        |



1 PWA analysis of AUGUSTIN 92 assigns  $0^{-+}$  quantum numbers to this state rather than  $1^{++}$  as before.

 $\eta(1295)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>55±5 OUR AVERAGE</b>				
57±23±21	20k	ADAMS	01B B852	18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
66±13	9082	MANAK	00A MPS	18 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
53±6		FUKUI	91C SPEC	8.95 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- |       |      |          |          |  |
|-------|------|----------|----------|--|
| <40   | 2100 | ALDE     | 97B GAM4 | 100 $\pi^- p \rightarrow \eta \pi^0 \pi^0 n$ |
| 44±20 | 2    | AUGUSTIN | 90 DM2   | $J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$ |
| ~ 70  |      | STANTON  | 79 CNTR  | 8.4 $\pi^- p \rightarrow n \eta 2\pi$        |
- 2 PWA analysis of AUGUSTIN 92 assigns  $0^{-+}$  quantum numbers to this state rather than  $1^{++}$  as before.

 $\eta(1295)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \eta \pi^+ \pi^-$	seen
$\Gamma_2 a_0(980) \pi$	seen
$\Gamma_3 \gamma \gamma$	
$\Gamma_4 \eta \pi^0 \pi^0$	seen
$\Gamma_5 \eta(\pi\pi)S\text{-wave}$	seen
$\Gamma_6 \sigma \eta$	
$\Gamma_7 K \bar{K} \pi$	

 $\eta(1295) \Gamma(i) \Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

$\Gamma(\eta \pi^+ \pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_1 \Gamma_3/\Gamma$
<b>&lt;0.066</b>	95		ACCIARRI	01G L3	183–202 $e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.6	90		AIHARA	88C TPC	$e^+ e^- \rightarrow e^+ e^- \eta \pi^+ \pi^-$	
<0.3			ANTREASYAN	87 CBAL	$e^+ e^- \rightarrow e^+ e^- \eta \pi \pi$	

 $\Gamma(K \bar{K} \pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ 

VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_7 \Gamma_3/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.014	90	3,4 AHOHE	05 CLE2	10.6 $e^+ e^- \rightarrow e^+ e^- K_S^0 K^\pm \pi^\mp$	

3 Using  $\eta(1295)$  mass and width 1294 MeV and 55 MeV, respectively.

4 Assuming three-body phase-space decay to  $K_S^0 K^\pm \pi^\mp$ .

 $\eta(1295)$  BRANCHING RATIOS

$\Gamma(a_0(980) \pi)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
not seen		BERTIN	97 OBLX	0.0 $\bar{p} p \rightarrow K^\pm (K^0) \pi^\mp \pi^+ \pi^-$	
seen		BIRMAN	88 MPS	8 $\pi^- p \rightarrow K^\pm \bar{K}^0 \pi^- n$	
large		ANDO	86 SPEC	8 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$	
large		STANTON	79 CNTR	8.4 $\pi^- p \rightarrow n \eta 2\pi$	

 $\Gamma(a_0(980) \pi)/\Gamma(\eta \pi^0 \pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_4$
<b>0.65±0.10</b>	5	ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta \pi^0 \pi^0 n$

5 Assuming that  $a_0(980)$  decays only to  $\eta \pi$ .

 $\Gamma(\eta(\pi\pi)S\text{-wave})/\Gamma(\eta \pi^0 \pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma_4$
<b>0.35±0.10</b>		ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta \pi^0 \pi^0 n$

 $\Gamma(a_0(980) \pi)/\Gamma(\sigma \eta)$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_6$
<b>0.48±0.22</b>	9082	MANAK	00A MPS	18 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$	

 $\eta(1295)$  REFERENCES

AHOHE	05	PR D71 072001	R. Ahohe <i>et al.</i>	(CLEO Collab.)
ACCIARRI	01G	PL B501 1	M. Acciarri <i>et al.</i>	(L3 Collab.)
ADAMS	01B	PL B516 264	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)
MANAK	00A	PR D62 012003	J.J. Manak <i>et al.</i>	(BNL E852 Collab.)
ALDE	97B	PAN 60 386	D. Alde <i>et al.</i>	(GAMS Collab.)
		Translated from YAF 60 458.		
BERTIN	97	PL B400 226	A. Bertin <i>et al.</i>	(OBLIX Collab.)
AUGUSTIN	92	PR D46 1951	J.E. Augustin, G. Cosme	(DM2 Collab.)
FUKUI	91C	PL B267 293	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
AUGUSTIN	90	PR D42 10	J.E. Augustin <i>et al.</i>	(DM2 Collab.)
AIHARA	88C	PR D38 1	H. Aihara <i>et al.</i>	(TPC2+ Collab.)
BIRMAN	88	PRL 61 1557	A. Birman <i>et al.</i>	(BNL FSU, IND, MASD)JP
ANTREASYAN	87	PR D36 2633	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
ANDO	86	PRL 57 1296	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)IJP
STANTON	79	PRL 42 346	N.R. Stanton <i>et al.</i>	(OSU, CARL, MCGI+)JP

 $\pi(1300)$ 

$$J^G(J^{PC}) = 1^-(0^{-+})$$

 $\pi(1300)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1300±100 OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1345±8±10	18k	1 SCHEGELSKY	06 RVUE	$\gamma \gamma \rightarrow \pi^+ \pi^- \pi^0$
1200±40	90k	SALVINI	04 OBLX	$\bar{p} p \rightarrow 2\pi^+ 2\pi^-$
1343±15±24		CHUNG	02 B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
1375±40		ABELE	01 CBAR	0.0 $\bar{p} d \rightarrow \pi^- 4\pi^0 p$
1275±15		BERTIN	97D OBLX	0.05 $\bar{p} p \rightarrow 2\pi^+ 2\pi^-$
~ 1114		ABELE	96 CBAR	0.0 $\bar{p} p \rightarrow 5\pi^0$
1190±30		ZIELINSKI	84 SPEC	200 $\pi^- Z \rightarrow Z 3\pi$
1240±30		BELLINI	82 SPEC	40 $\pi^- A \rightarrow A 3\pi$
1273±50		2 AARON	81 RVUE	
1342±20		BONESINI	81 OMEG	12 $\pi^- p \rightarrow p 3\pi$
~ 1400		DAUM	81B SPEC	63,94 $\pi^- p$

1 From analysis of L3 data at 183–209 GeV.

2 Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from DAUM 80 and DANKOWYCH 81.

 $\pi(1300)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>200 to 600 OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
260±20±30	18k	3 SCHEGELSKY	06 RVUE	$\gamma \gamma \rightarrow \pi^+ \pi^- \pi^0$
470±120	90k	SALVINI	04 OBLX	$\bar{p} p \rightarrow 2\pi^+ 2\pi^-$
449±39±47		CHUNG	02 B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
268±50		ABELE	01 CBAR	0.0 $\bar{p} d \rightarrow \pi^- 4\pi^0 p$
218±100		BERTIN	97D OBLX	0.05 $\bar{p} p \rightarrow 2\pi^+ 2\pi^-$
~ 340		ABELE	96 CBAR	0.0 $\bar{p} p \rightarrow 5\pi^0$

# Meson Particle Listings

## $\pi(1300)$ , $a_2(1320)$

440± 80	ZIELINSKI	84	SPEC	200 $\pi^+ Z \rightarrow Z 3\pi$
360±120	BELLINI	82	SPEC	40 $\pi^- A \rightarrow A 3\pi$
580±100	<sup>4</sup> AARON	81	RVUE	
220± 70	BONESINI	81	OMEG	12 $\pi^- p \rightarrow p 3\pi$
~ 600	DAUM	81B	SPEC	63,94 $\pi^- p$

<sup>3</sup>From analysis of L3 data at 183–209 GeV.

<sup>4</sup>Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from DAUM 80 and DANKOWYCH 81.

### $\pi(1300)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\rho\pi$	seen
$\Gamma_2$ $\pi(\pi\pi)s$ -wave	seen
$\Gamma_3$ $\gamma\gamma$	

### $\pi(1300)$ $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(\rho\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_1\Gamma_3/\Gamma$
<0.085	90	ACCIARRI	97T	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.8	95	<sup>5</sup> SCHEGELSKY	06	RVUE $\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$	
<0.54	90	ALBRECHT	97B	ARG $e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	

<sup>5</sup>From analysis of L3 data at 183–209 GeV.

### $\pi(1300)$ BRANCHING RATIOS

$\Gamma(\pi(\pi\pi)s\text{-wave})/\Gamma(\rho\pi)$	CL%	EVTs	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_1$
2.2 ± 0.4	90k		SALVINI	04	OBLX $\bar{p}p \rightarrow 2\pi^+2\pi^-$	
seen			CHUNG	02	B852 $18.3\pi^-p \rightarrow \pi^+2\pi^-\pi^0$	
<0.15	90		ABELE	01	CBAR $0.0\bar{p}d \rightarrow \pi^-4\pi^0p$	
2.12			<sup>6</sup> AARON	81	RVUE	

<sup>6</sup>Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from DAUM 80 and DANKOWYCH 81.

### $\pi(1300)$ REFERENCES

SCHEGELSKY	06	EPJ A27 199	V.A. Schegelsky <i>et al.</i>	(OBELIX Collab.)
SALVINI	04	EPJ C35 21	P. Salvini <i>et al.</i>	(S.U. Chung <i>et al.</i> )
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
ABELE	01	EPJ C19 667	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACCIARRI	97T	PL B413 147	M. Acciari <i>et al.</i>	(L3 Collab.)
ALBRECHT	97B	ZPHY C74 469	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BERTIN	97D	PL B414 220	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ABELE	96	PL B380 453	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ZIELINSKI	84	PR D30 1855	M. Zielinski <i>et al.</i>	(ROCH, MINN, FNAL)
BELLINI	82	PRL 48 1697	G. Bellini <i>et al.</i>	(MILA, BGNA, JINR)
AARON	81	PR D24 1207	R.A. Aaron, R.S. Longacre	(NEAS, BNL)
BONESINI	81	PL 103B 75	M. Bonesini <i>et al.</i>	(MILA, LIPV, DARE+)
DANKOWYCH...	81	PRL 46 580	J.A. Dankowycz <i>et al.</i>	(TNT0, BNL, CARL+)
DAUM	81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
DAUM	80	PL B9B 281	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
BOWLER	75	NP B97 227	M.G. Bowler <i>et al.</i>	(OXFTP, DARE)

## $a_2(1320)$

$$I^G(J^{PC}) = 1^-(2^{++})$$

### $a_2(1320)$ MASS

VALUE (MeV)	DOCUMENT ID
<b>1318.3<sup>+0.5</sup><sub>-0.6</sub> OUR AVERAGE</b>	Includes data from the 4 datablocks that follow this one. Error includes scale factor of 1.2.

### 3 $\pi$ MODE

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

<b>1319.0<sup>+1.0</sup><sub>-1.3</sub> OUR AVERAGE</b>	Error includes scale factor of 1.4. See the ideogram below.
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1321 ± 1 <sup>+0</sup> <sub>-7</sub>	420k	ALEKSEEV	10	COMP	190 $\pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
1326 ± 2 ± 2		CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
1317 ± 3		BARBERIS	98B		450 $p p \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
1323 ± 4 ± 3		ACCIARRI	97T	L3	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$
1320 ± 7		ALBRECHT	97B	ARG	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$
1311.3± 1.6± 3.0	72.4k	AMELIN	96	VES	36 $\pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1310 ± 5		ARMSTRONG	90	OMEG	0 300.0 $p p \rightarrow p p \pi^+ \pi^- \pi^0$

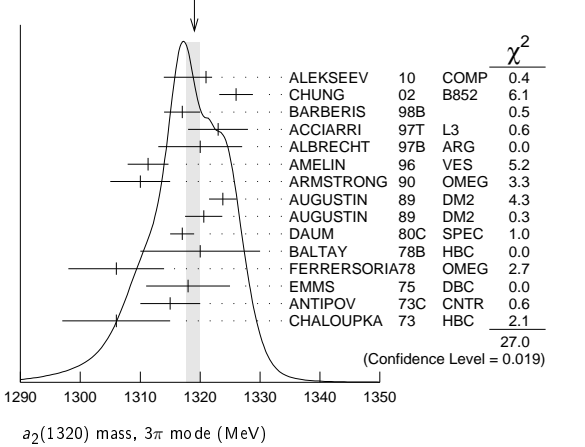
1323.8± 2.3	4022	AUGUSTIN	89	DM2	± $J/\psi \rightarrow \rho^\pm a_2^\mp$
1320.6± 3.1	3562	AUGUSTIN	89	DM2	0 $J/\psi \rightarrow \rho^0 a_2^0$
1317 ± 2	25k	<sup>1</sup> DAUM	80c	SPEC	− 63,94 $\pi^- p \rightarrow 3\pi p$
1320 ± 10	1097	<sup>1</sup> BALTAY	78B	HBC	+0 15 $\pi^+ p \rightarrow p 4\pi$
1306 ± 8		FERRERSORIA	78	OMEG	− 9 $\pi^- p \rightarrow p 3\pi$
1318 ± 7	1.6k	<sup>1</sup> EMMS	75	DBC	0 4 $\pi^+ n \rightarrow p(3\pi)^0$
1315 ± 5		<sup>1</sup> ANTIPOV	73c	CNTR	− 25,40 $\pi^- p \rightarrow \rho \eta \pi^-$

1306 ± 9	1580	CHALOUKPA	73	HBC	− 3.9 $\pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1300 ± 2 ± 4	18k	<sup>2</sup> SCHEGELSKY	06	RVUE	0 $\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
1305 ± 14		CONDO	93	SHF	$\gamma p \rightarrow n\pi^+\pi^-\pi^-$
1310 ± 2		<sup>1</sup> EVANGELIS...	81	OMEG	− 12 $\pi^- p \rightarrow 3\pi p$
1343 ± 11	490	BALTAY	78B	HBC	0 15 $\pi^+ p \rightarrow \Delta 3\pi$
1309 ± 5	5k	BINNIE	71	MMS	− $\pi^- p$ near $a_2$ thresh- old
1299 ± 6	28k	BOWEN	71	MMS	− 5 $\pi^- p$
1300 ± 6	24k	BOWEN	71	MMS	+ 5 $\pi^+ p$
1309 ± 4	17k	BOWEN	71	MMS	− 7 $\pi^- p$
1306 ± 4	941	ALSTON...	70	HBC	+ 7.0 $\pi^+ p \rightarrow 3\pi p$

<sup>1</sup>From a fit to  $J^P = 2^+ \rho\pi$  partial wave.

<sup>2</sup>From analysis of L3 data at 183–209 GeV.

WEIGHTED AVERAGE  
1319.0+1.0-1.3 (Error scaled by 1.4)



### $K\bar{K}$ MODE

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

### 1318.1± 0.7 OUR AVERAGE

1319 ± 5	4700	<sup>1,2</sup> CLELAND	82B	SPEC	+ 50 $\pi^+ p \rightarrow K_S^0 K^+ p$
1324 ± 6	5200	<sup>1,2</sup> CLELAND	82B	SPEC	− 50 $\pi^- p \rightarrow K_S^0 K^- p$
1320 ± 2	4000	CHABAUD	80	SPEC	− 17 $\pi^- A \rightarrow K_S^0 K^- A$
1312 ± 4	11000	CHABAUD	78	SPEC	− 9.8 $\pi^- p \rightarrow K^- K_S^0 p$
1316 ± 2	4730	CHABAUD	78	SPEC	− 18.8 $\pi^- p \rightarrow K^- K_S^0 p$
1318 ± 1		<sup>1,3</sup> MARTIN	78D	SPEC	− 10 $\pi^- p \rightarrow K_S^0 K^- p$
1320 ± 2	2724	MARGULIE	76	SPEC	− 23 $\pi^- p \rightarrow K^- K_S^0 p$
1313 ± 4	730	FOLEY	72	CNTR	− 20.3 $\pi^- p \rightarrow K^- K_S^0 p$
1319 ± 3	1500	<sup>3</sup> GRAY	71	ASPK	− 17.2 $\pi^+ p \rightarrow K^- K_S^0 p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1304 ± 10	870	<sup>4</sup> SCHEGELSKY	06A	RVUE	0 $\gamma\gamma \rightarrow K_S^0 K_S^0$
1330 ± 11	1000	<sup>1,2</sup> CLELAND	82B	SPEC	+ 30 $\pi^+ p \rightarrow K_S^0 K^+ p$
1324 ± 5	350	HYAMS	78	ASPK	+ 12.7 $\pi^+ p \rightarrow K^+ K_S^0 p$

<sup>1</sup>From a fit to  $J^P = 2^+$  partial wave.

<sup>2</sup>Number of events evaluated by us.

<sup>3</sup>Systematic error in mass scale subtracted.

<sup>4</sup>From analysis of L3 data at 91 and 183–209 GeV.

### $\eta\pi$ MODE

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

### 1317.7± 1.4 OUR AVERAGE

1308 ± 9		BARBERIS	00H		450 $p p \rightarrow p_f \eta \pi^0 p_s$
1316 ± 9		BARBERIS	00H		450 $p p \rightarrow \Delta_f^{++} \eta \pi^- p_s$
1317 ± 1 ± 2		THOMPSON	97	MPS	18 $\pi^- p \rightarrow \eta \pi^- p$
1315 ± 5 ± 2		<sup>1</sup> AMSLER	94D	CBAR	0.0 $\bar{p} p \rightarrow \pi^0 \pi^0 \eta$
1325.1± 5.1		AOYAGI	93	BKEI	$\pi^- p \rightarrow \eta \pi^- p$
1317.7± 1.4± 2.0		BELADIDZE	93	VES	37 $\pi^- N \rightarrow \eta \pi^- N$
1323 ± 8	1000	<sup>2</sup> KEY	73	OSPK	− 6 $\pi^- p \rightarrow p \pi^- \eta$

See key on page 601

## Meson Particle Listings

 $a_2(1320)$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

1315 ±12		<sup>3</sup> ADOLPH	15	COMP	191 $\pi^- p \rightarrow \eta^{(\prime)} \pi^- p$
1309 ± 4		ANISOVICH	09	RVUE	$\bar{p} p, \pi N$
1324 ± 5		ARMSTRONG	93c	E760	0 $\bar{p} p \rightarrow \pi^0 \eta \eta \rightarrow 6\gamma$
1336.2 ± 1.7	2561	DELFOSE	81	SPEC +	$\pi^\pm p \rightarrow p \pi^\pm \eta$
1330.7 ± 2.4	1653	DELFOSE	81	SPEC -	$\pi^\pm p \rightarrow p \pi^\pm \eta$
1324 ± 8	6200	<sup>2,4</sup> CONFORTO	73	OSPK -	$6 \pi^- p \rightarrow p \text{MM}^-$

<sup>1</sup> The systematic error of 2 MeV corresponds to the spread of solutions.<sup>2</sup> Error includes 5 MeV systematic mass-scale error.<sup>3</sup> ADOLPH 15 value is derived from a Breit-Wigner fit with mass-dependent width taking the  $\eta \pi$  and  $\rho \pi$  channels into account.<sup>4</sup> Missing mass with enriched MMS =  $\eta \pi^-$ ,  $\eta = 2\gamma$ . $\eta/\pi$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
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The data in this block is included in the average printed for a previous datablock.

**1322 ± 7 OUR AVERAGE**

1318 ± 8 $^{+3}_{-5}$	IVANOV	01	B852	18 $\pi^- p \rightarrow \eta' \pi^- p$
1327.0 ± 10.7	BELADIDZE	93	VES	37 $\pi^- N \rightarrow \eta' \pi^- N$

 $a_2(1320)$  WIDTH**3 $\pi$  MODE**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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**105.0 $^{+1.6}_{-1.9}$  OUR AVERAGE**

110 ± 2 $^{+2}_{-15}$	420k	ALEKSEEV	10	COMP	190 $\pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
108 ± 3 ± 15		CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
120 ± 10		BARBERIS	98b		450 $p p \rightarrow \rho_f \pi^+ \pi^- \pi^0 p_s$
105 ± 10 ± 11		ACCIARRI	97t	L3	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^- \pi^0$
120 ± 10		ALBRECHT	97b	ARG	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^- \pi^0$
103.0 ± 6.0 ± 3.3	72.4k	AMELIN	96	VES	36 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- n$
120 ± 10		ARMSTRONG	90	OMEG 0	300.0 $p p \rightarrow \rho \rho \pi^+ \pi^- \pi^0$
107.0 ± 9.7	4022	AUGUSTIN	89	DM2 ±	$J/\psi \rightarrow \rho^\pm a_2^\mp$
118.5 ± 12.5	3562	AUGUSTIN	89	DM2 0	$J/\psi \rightarrow \rho^0 a_2^0$
97 ± 5		<sup>1</sup> EVANGELIS...	81	OMEG -	12 $\pi^- p \rightarrow 3\pi p$
96 ± 9	25k	<sup>1</sup> DAUM	80c	SPEC -	63.94 $\pi^- p \rightarrow 3\pi p$
110 ± 15	1097	<sup>1</sup> BALTAY	78b	HBC +0	15 $\pi^+ p \rightarrow p 4\pi$
112 ± 18	1.6k	<sup>1</sup> EMMS	75	DBC 0	4 $\pi^+ n \rightarrow p(3\pi)^0$
122 ± 14	1.2k	<sup>1,2</sup> WAGNER	75	HBC 0	7 $\pi^+ p \rightarrow \Delta^{++}(3\pi)^0$

115 ± 15		<sup>1</sup> ANTIPOV	73c	CNTR -	25.40 $\pi^- p \rightarrow \rho \eta \pi^-$
99 ± 15	1580	CHALOUPKA	73	HBC -	3.9 $\pi^- p$
105 ± 5	28k	BOWEN	71	MMS -	5 $\pi^- p$
99 ± 5	24k	BOWEN	71	MMS +	5 $\pi^+ p$
103 ± 5	17k	BOWEN	71	MMS -	7 $\pi^- p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

117 ± 6 ± 20	18k	<sup>3</sup> SCHEGELSKY	06	RVUE 0	$\gamma \gamma \rightarrow \pi^+ \pi^- \pi^0$
120 ± 40		CONDO	93	SHF	$\gamma p \rightarrow n \pi^+ \pi^+ \pi^-$
115 ± 14	490	BALTAY	78b	HBC 0	15 $\pi^+ p \rightarrow \Delta 3\pi$
72 ± 16	5k	BINNIE	71	MMS -	$\pi^- p$ near $a_2$ thresh-old
79 ± 12	941	ALSTON-...	70	HBC +	7.0 $\pi^+ p \rightarrow 3\pi p$

<sup>1</sup> From a fit to  $J^P = 2^+ \rho \pi$  partial wave.<sup>2</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.<sup>3</sup> From analysis of L3 data at 183–209 GeV. **$K\bar{K}$  AND  $\eta \pi$  MODES**

VALUE (MeV)	DOCUMENT ID
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**107 ± 5 OUR ESTIMATE****110.4 ± 1.7 OUR AVERAGE** Includes data from the 2 datablocks that follow this one. **$K\bar{K}$  MODE**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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The data in this block is included in the average printed for a previous datablock.

**109.8 ± 2.4 OUR AVERAGE**

112 ± 20	4700	<sup>1,2</sup> CLELAND	82b	SPEC +	50 $\pi^+ p \rightarrow K^0 K^+ p$
120 ± 25	5200	<sup>1,2</sup> CLELAND	82b	SPEC -	50 $\pi^- p \rightarrow K^0 K^- p$
106 ± 4	4000	CHABAUD	80	SPEC -	17 $\pi^- A \rightarrow K_S^0 K^- A$
126 ± 11	11000	CHABAUD	78	SPEC -	9.8 $\pi^- p \rightarrow K^- K_S^0 p$
101 ± 8	4730	CHABAUD	78	SPEC -	18.8 $\pi^- p \rightarrow K^- K_S^0 p$
113 ± 4		<sup>1,3</sup> MARTIN	78d	SPEC -	10 $\pi^- p \rightarrow K_S^0 K^- p$
105 ± 8	2724	<sup>3</sup> MARGULIE	76	SPEC -	23 $\pi^- p \rightarrow K^- K_S^0 p$
113 ± 19	730	FOLEY	72	CNTR -	20.3 $\pi^- p \rightarrow K^- K_S^0 p$
123 ± 13	1500	<sup>3</sup> GRAYER	71	ASPK -	17.2 $\pi^- p \rightarrow K^- K_S^0 p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

120 ± 15	870	<sup>4</sup> SCHEGELSKY	06a	RVUE 0	$\gamma \gamma \rightarrow K_S^0 K_S^0$
121 ± 51	1000	<sup>1,2</sup> CLELAND	82b	SPEC +	30 $\pi^+ p \rightarrow K_S^0 K^+ p$
110 ± 18	350	HYAMS	78	ASPK +	12.7 $\pi^+ p \rightarrow K^+ K_S^0 p$

<sup>1</sup> From a fit to  $J^P = 2^+$  partial wave.<sup>2</sup> Number of events evaluated by us.<sup>3</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.<sup>4</sup> From analysis of L3 data at 91 and 183–209 GeV. **$\eta \pi$  MODE**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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The data in this block is included in the average printed for a previous datablock.

**111.1 ± 2.4 OUR AVERAGE**

115 ± 20		BARBERIS	00h		450 $p p \rightarrow p_f \eta \pi^0 p_s$
112 ± 14		BARBERIS	00h		450 $p p \rightarrow \Delta_f^{++} \eta \pi^- p_s$
112 ± 3 ± 2		<sup>1</sup> AMSLER	94d	CBAR	0.0 $\bar{p} p \rightarrow \pi^0 \pi^0 \eta$
103 ± 6 ± 3		BELADIDZE	93	VES	37 $\pi^- N \rightarrow \eta \pi^- N$
112.2 ± 5.7	2561	DELFOSE	81	SPEC +	$\pi^\pm p \rightarrow p \pi^\pm \eta$
116.6 ± 7.7	1653	DELFOSE	81	SPEC -	$\pi^\pm p \rightarrow p \pi^\pm \eta$
108 ± 9	1000	KEY	73	OSPK -	6 $\pi^- p \rightarrow p \pi^- \eta$

• • • We do not use the following data for averages, fits, limits, etc. • • •

119 ± 14		<sup>2</sup> ADOLPH	15	COMP	191 $\pi^- p \rightarrow \eta^{(\prime)} \pi^- p$
110 ± 4		ANISOVICH	09	RVUE	$\bar{p} p, \pi N$
127 ± 2 ± 2		<sup>3</sup> THOMPSON	97	MPS	18 $\pi^- p \rightarrow \eta \pi^- p$
118 ± 10		ARMSTRONG	93c	E760	0 $\bar{p} p \rightarrow \pi^0 \eta \eta \rightarrow 6\gamma$
104 ± 9	6200	<sup>4</sup> CONFORTO	73	OSPK -	6 $\pi^- p \rightarrow p \text{MM}^-$

<sup>1</sup> The systematic error of 2 MeV corresponds to the spread of solutions.<sup>2</sup> ADOLPH 15 value is derived from a Breit-Wigner fit with mass-dependent width taking the  $\eta \pi$  and  $\rho \pi$  channels into account.<sup>3</sup> Resolution is not unfolded.<sup>4</sup> Missing mass with enriched MMS =  $\eta \pi^-$ ,  $\eta = 2\gamma$ . **$\eta/\pi$  MODE**

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
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**119 ± 25 OUR AVERAGE**

140 ± 35 ± 20	IVANOV	01	B852	18 $\pi^- p \rightarrow \eta' \pi^- p$
106 ± 32	BELADIDZE	93	VES	37 $\pi^- N \rightarrow \eta' \pi^- N$

 $a_2(1320)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ 3 $\pi$	(70.1 ± 2.7 ) %	S=1.2
$\Gamma_2$ $\rho(770) \pi$		
$\Gamma_3$ $f_2(1270) \pi$		
$\Gamma_4$ $\rho(1450) \pi$		
$\Gamma_5$ $\eta \pi$	(14.5 ± 1.2 ) %	
$\Gamma_6$ $\omega \pi \pi$	(10.6 ± 3.2 ) %	S=1.3
$\Gamma_7$ $K \bar{K}$	( 4.9 ± 0.8 ) %	
$\Gamma_8$ $\eta'(958) \pi$	( 5.5 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_9$ $\pi^\pm \gamma$	( 2.91 ± 0.27 ) × 10 <sup>-3</sup>	
$\Gamma_{10}$ $\gamma \gamma$	( 9.4 ± 0.7 ) × 10 <sup>-6</sup>	
$\Gamma_{11}$ $e^+ e^-$	< 5 × 10 <sup>-9</sup>	CL=90%

## CONSTRAINED FIT INFORMATION

An overall fit to 5 branching ratios uses 18 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 9.3$  for 15 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_5$	10		
$x_6$	-89	-46	
$x_7$	-1	-2	-24
	$x_1$	$x_5$	$x_6$

 $a_2(1320)$  PARTIAL WIDTHS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

18.5 ± 3.0	870	<sup>1</sup> SCHEGELSKY	06a	RVUE 0	$\gamma \gamma \rightarrow K_S^0 K_S^0$
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<sup>1</sup> From analysis of L3 data at 91 and 183–209 GeV, using  $\Gamma(a_2(1320) \rightarrow \gamma \gamma) = 0.91$  keV and SU(3) relations.

# Meson Particle Listings

## $a_2(1320)$

### $\Gamma(K\bar{K})$ $\Gamma_7$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$7.0^{+2.0}_{-1.5}$	870	<sup>1</sup> SCHEGELSKY 06A	RVUE	0	$\gamma\gamma \rightarrow K_S^0 K_S^0$
<sup>1</sup> From analysis of L3 data at 91 and 183–209 GeV, using $\Gamma(a_2(1320) \rightarrow \gamma\gamma) = 0.91$ keV and SU(3) relations.					

### $\Gamma(\pi^\pm\gamma)$ $\Gamma_9$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>311 ± 25 OUR AVERAGE</b>					
$358 \pm 6 \pm 42$		<sup>1</sup> ADOLPH 14	COMP	—	$190 \pi^- \text{Pb} \rightarrow \pi^+ \pi^- \pi^- \text{Pb}'$
$284 \pm 25 \pm 25$	7.1k	MOLCHANOV 01	SELX		$600 \pi^- A \rightarrow \pi^+ \pi^- \pi^- A$
$295 \pm 60$		CIHANGIR 82	SPEC	+	$200 \pi^+ A$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$461 \pm 110$		<sup>2</sup> MAY 77	SPEC	±	$9.7 \gamma A$
<sup>1</sup> Primakoff reaction using $a_2(1320) \rightarrow 3\pi$ branching ratio of 70.1%.					
<sup>2</sup> Assuming one-pion exchange.					

### $\Gamma(\gamma\gamma)$ $\Gamma_{10}$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.00 ± 0.06 OUR AVERAGE</b>					
$0.98 \pm 0.05 \pm 0.09$		ACCARI 97t	L3		$e^+e^- \rightarrow e^+e^- \pi^+ \pi^- \pi^0$
$0.96 \pm 0.03 \pm 0.13$		ALBRECHT 97b	ARG		$e^+e^- \rightarrow e^+e^- \pi^+ \pi^- \pi^0$
$1.26 \pm 0.26 \pm 0.18$	36	BARU 90	MD1		$e^+e^- \rightarrow e^+e^- \pi^+ \pi^- \pi^0$
$1.00 \pm 0.07 \pm 0.15$	415	BEHREND 90c	CELL	0	$e^+e^- \rightarrow e^+e^- \pi^+ \pi^- \pi^0$
$1.03 \pm 0.13 \pm 0.21$		BUTLER 90	MRK2		$e^+e^- \rightarrow e^+e^- \pi^+ \pi^- \pi^0$
$1.01 \pm 0.14 \pm 0.22$	85	OEST 90	JADE		$e^+e^- \rightarrow e^+e^- \pi^0 \eta$
$0.90 \pm 0.27 \pm 0.15$	56	<sup>1</sup> ALTHOFF 86	TASS	0	$e^+e^- \rightarrow e^+e^- 3\pi$
$1.14 \pm 0.20 \pm 0.26$		<sup>2</sup> ANTREASIAN 86	CBAL	0	$e^+e^- \rightarrow e^+e^- \pi^0 \eta$
$1.06 \pm 0.18 \pm 0.19$		BERGER 84c	PLUT	0	$e^+e^- \rightarrow e^+e^- 3\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.81 \pm 0.19 \pm 0.42$	35	<sup>1</sup> BEHREND 83b	CELL	0	$e^+e^- \rightarrow e^+e^- 3\pi$
$0.77 \pm 0.18 \pm 0.27$	22	<sup>2</sup> EDWARDS 82f	CBAL	0	$e^+e^- \rightarrow e^+e^- \pi^0 \eta$
<sup>1</sup> From $\rho\pi$ decay mode.					
<sup>2</sup> From $\eta\pi^0$ decay mode.					

### $\Gamma(e^+e^-)$ $\Gamma_{11}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.56</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<25	90	VOROBYEV 88	ND	$e^+e^- \rightarrow \pi^0 \eta$

### $a_2(1320) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

### $\Gamma(3\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ $\Gamma_1\Gamma_{10}/\Gamma$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.65 \pm 0.02 \pm 0.02$	18k	<sup>1</sup> SCHEGELSKY 06	RVUE	$\gamma\gamma \rightarrow \pi^+ \pi^- \pi^0$
<sup>1</sup> From analysis of L3 data at 183–209 GeV.				

### $\Gamma(\eta\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ $\Gamma_5\Gamma_{10}/\Gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.145^{+0.097}_{-0.034}$	<sup>1</sup> UEHARA 09A	BELL	$e^+e^- \rightarrow e^+e^- \eta\pi^0$
<sup>1</sup> From the $D_2$ -wave. The fraction of the $D_0$ -wave is $3.4^{+2.3}_{-1.1}\%$ .			

### $\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ $\Gamma_7\Gamma_{10}/\Gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>0.126 ± 0.007 ± 0.028</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.081 \pm 0.006 \pm 0.027$	<sup>2</sup> ALBRECHT 90G	ARG	$e^+e^- \rightarrow e^+e^- K^+ K^-$
<sup>1</sup> Using an incoherent background.			
<sup>2</sup> Using a coherent background.			

### $a_2(1320)$ BRANCHING RATIOS

### $[\Gamma(f_2(1270)\pi) + \Gamma(\rho(1450)\pi)]/\Gamma(\rho(770)\pi)$ $(\Gamma_3+\Gamma_4)/\Gamma_2$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;0.12</b>	90	ABRAMOVI... 70b	HBC	—	$3.93 \pi^- p$

### $\Gamma(\eta\pi)/\Gamma(3\pi)$ $\Gamma_5/\Gamma_1$

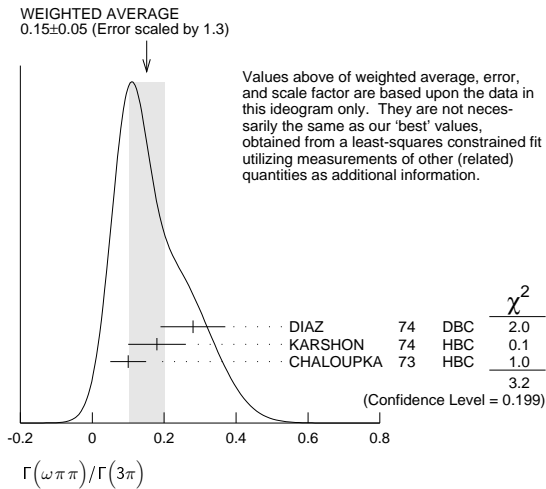
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.207 ± 0.018 OUR FIT</b>					
<b>0.213 ± 0.020 OUR AVERAGE</b>					
$0.18 \pm 0.05$		FORINO 76	HBC		$11 \pi^- p$
$0.22 \pm 0.05$	52	ANTIPOV 73	CNTR	—	$40 \pi^- p$
$0.211 \pm 0.044$	149	CHALOUPKA 73	HBC	—	$3.9 \pi^- p$
$0.246 \pm 0.042$	167	ALSTON-... 71	HBC	+	$7.0 \pi^+ p$
$0.25 \pm 0.09$	15	BOECKMANN 70	HBC	+	$5.0 \pi^+ p$
$0.23 \pm 0.08$	22	ASCOLI 68	HBC	—	$5 \pi^- p$
$0.12 \pm 0.08$		CHUNG 68	HBC	—	$3.2 \pi^- p$
$0.22 \pm 0.09$		CONTE 67	HBC	—	$11.0 \pi^- p$

### $\Gamma(\omega\pi\pi)/\Gamma(3\pi)$ $\Gamma_6/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.15 ± 0.05 OUR FIT</b>					
<b>0.15 ± 0.05 OUR AVERAGE</b>					
Error includes scale factor of 1.3.					
$0.28 \pm 0.09$	60	DIAZ 74	DBC	0	$6 \pi^+ n$
$0.18 \pm 0.08$		<sup>1</sup> KARSHON 74	HBC		Avg. of above two
$0.10 \pm 0.05$	279	<sup>2</sup> CHALOUPKA 73	HBC	—	$3.9 \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.29 \pm 0.08$	140	<sup>1</sup> KARSHON 74	HBC	0	$4.9 \pi^+ p$
$0.10 \pm 0.04$	60	<sup>1</sup> KARSHON 74	HBC	+	$4.9 \pi^- p$
$0.19 \pm 0.08$		DEFOIX 73	HBC	0	$0.7 \bar{p} p$

<sup>1</sup> KARSHON 74 suggest an additional  $l = 0$  state strongly coupled to  $\omega\pi\pi$  which could explain discrepancies in branching ratios and masses. We use a central value and a systematic spread.

<sup>2</sup> Decays to  $b_1(1040)\pi$ ,  $b_1 \rightarrow \omega\pi$ . Error increased to account for possible systematic errors of complicated analysis.



### $\Gamma(K\bar{K})/\Gamma(3\pi)$ $\Gamma_7/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.070 ± 0.012 OUR FIT</b>					
<b>0.078 ± 0.017</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.011 \pm 0.003$		CHABAUD 78	RVUE		
$0.011 \pm 0.003$		<sup>1</sup> BERTIN 98b	OBLX	—	$0.0 \bar{p} p \rightarrow K^\pm K_S \pi^\mp$
$0.056 \pm 0.014$	50	<sup>2</sup> CHALOUPKA 73	HBC	—	$3.9 \pi^- p$
$0.097 \pm 0.018$	113	<sup>2</sup> ALSTON-... 71	HBC	+	$7.0 \pi^+ p$
$0.06 \pm 0.03$		<sup>2</sup> ABRAMOVI... 70b	HBC	—	$3.93 \pi^- p$
$0.054 \pm 0.022$		<sup>2</sup> CHUNG 68	HBC	—	$3.2 \pi^- p$
<sup>1</sup> Using 4 $\pi$ data from BERTIN 97D.					
<sup>2</sup> Included in CHABAUD 78 review.					

### $\Gamma(K\bar{K})/\Gamma(\eta\pi)$ $\Gamma_7/\Gamma_5$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.08 \pm 0.02$	<sup>1</sup> BERTIN 98b	OBLX	$0.0 \bar{p} p \rightarrow K^\pm K_S \pi^\mp$
<sup>1</sup> Using $\eta\pi\pi$ data from AMSLER 94D.			

### $\Gamma(\eta\pi)/[\Gamma(3\pi) + \Gamma(\eta\pi) + \Gamma(K\bar{K})]$ $\Gamma_5/(\Gamma_1+\Gamma_5+\Gamma_7)$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.162 ± 0.012 OUR FIT</b>					
<b>0.140 ± 0.028 OUR AVERAGE</b>					
$0.13 \pm 0.04$		ESPIGAT 72	HBC	±	$0.0 \bar{p} p$
$0.15 \pm 0.04$	34	BARNHAM 71	HBC	+	$3.7 \pi^+ p$

See key on page 601

## Meson Particle Listings

 $a_2(1320)$ ,  $f_0(1370)$ 

$\Gamma(K\bar{K})/[\Gamma(3\pi) + \Gamma(\eta\pi) + \Gamma(K\bar{K})]$					$\Gamma_7/(\Gamma_1 + \Gamma_5 + \Gamma_7)$				
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT				
<b>0.054 ± 0.009 OUR FIT</b>									
<b>0.048 ± 0.012 OUR AVERAGE</b>									
0.05 ± 0.02		TOET	73	HBC	+	5 $\pi^+ p$			
0.09 ± 0.04		TOET	73	HBC	0	5 $\pi^+ p$			
0.03 ± 0.02	8	<sup>1</sup> DAMERI	72	HBC	—	11 $\pi^- p$			
0.06 ± 0.03	17	BARNHAM	71	HBC	+	3.7 $\pi^+ p$			
• • • We do not use the following data for averages, fits, limits, etc. • • •									
0.020 ± 0.004		<sup>2</sup> ESPIGAT	72	HBC	±	0.0 $\bar{p}p$			

<sup>1</sup> Montanet agrees. Vlada.<sup>2</sup> Not averaged because of discrepancy between masses from  $K\bar{K}$  and  $\rho\pi$  modes.

$\Gamma(\eta'(958)\pi)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$				
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<0.006	95	ALDE	92b	GAM2		38,100 $\pi^- p \rightarrow \eta' \pi^0 n$			
<0.02	97	BARNHAM	71	HBC	+	3.7 $\pi^+ p$			
0.004 ± 0.004		<sup>1</sup> BOESEBECK	68	HBC	+	8 $\pi^+ p$			
<sup>1</sup> No longer valid since $\Gamma(K\bar{K})/\Gamma(3\pi)$ value has changed (MORRISON 71).									

$\Gamma(\eta'(958)\pi)/\Gamma(3\pi)$					$\Gamma_8/\Gamma_1$				
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<0.011	90	EISENSTEIN	73	HBC	—	5 $\pi^- p$			
<0.04		ALSTON-...	71	HBC	+	7.0 $\pi^+ p$			
0.04 ± 0.03 — 0.04		BOECKMANN	70	HBC	0	5.0 $\pi^+ p$			

$\Gamma(\eta'(958)\pi)/\Gamma(\eta\pi)$					$\Gamma_8/\Gamma_5$				
VALUE	DOCUMENT ID	TECN	COMMENT						
<b>0.038 ± 0.005 OUR AVERAGE</b>									
0.05 ± 0.02	ADOLPH	15	COMP	191 $\pi^- p \rightarrow \eta^{(\prime)} \pi^- p$					
0.032 ± 0.009	ABELE	97c	CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \eta'$					
0.047 ± 0.010 ± 0.004	<sup>1</sup> BELADIDZE	93	VES	37 $\pi^- N \rightarrow \bar{a}_2^- N$					
0.034 ± 0.008 ± 0.005	BELADIDZE	92	VES	36 $\pi^- C \rightarrow \bar{a}_2^- C$					
<sup>1</sup> Using $B(\eta' \rightarrow \pi^+ \pi^- \eta) = 0.441$ , $B(\eta \rightarrow \gamma\gamma) = 0.389$ and $B(\eta \rightarrow \pi^+ \pi^- \pi^0) = 0.236$ .									

$\Gamma(\pi^\pm \gamma)/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$				
VALUE	DOCUMENT ID	TECN	COMMENT						
• • • We do not use the following data for averages, fits, limits, etc. • • •									
0.005 ± 0.005 — 0.003	<sup>1</sup> EISENBERG	72	HBC	4.3, 5.25, 7.5 $\gamma p$					
<sup>1</sup> Pion-exchange model used in this estimation.									

$\Gamma(e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$				
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID	TECN	COMMENT					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<6	90	ACHASOV	00k	SND	$e^+ e^- \rightarrow \pi^0 \pi^0$				

 $a_2(1320)$  REFERENCES

ADOLPH	15	PL B740 303	M. Adolph <i>et al.</i>	(COMPASS Collab.)
ADOLPH	14	EPJ A50 79	C. Adolph <i>et al.</i>	(COMPASS Collab.)
ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i>	(COMPASS Collab.)
ANISOVICH	09	UMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
UEHARA	09A	PR D80 032001	S. Uehara <i>et al.</i>	(BELLE Collab.)
SCHEGELSKY	06	EPJ A27 199	V.A. Schegelsky <i>et al.</i>	
SCHEGELSKY	06A	EPJ A27 207	V.A. Schegelsky <i>et al.</i>	
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
INOVANOV	01	PRL 86 3977	E.I. Ivanov <i>et al.</i>	(BNL E852 Collab.)
MOLCHANOV	01	PL B521 171	V.V. Molchanov <i>et al.</i>	(FNAL SELEX Collab.)
ACHASOV	00K	PL B492 8	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
BARBERIS	00H	PL B488 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	98B	PL B422 399	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BERTIN	98B	PL B434 180	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ABELE	97C	PL B404 179	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACCARIARI	97T	PL B413 147	M. Acciarri <i>et al.</i>	(L3 Collab.)
ALBRECHT	97B	ZPHY C74 469	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
THOMPSON	97	PRL 79 1630	D.R. Thompson <i>et al.</i>	(BNL E852 Collab.)
AMELIN	96	ZPHY C70 71	D.V. Amelin <i>et al.</i>	(SERP, T-BIL Collab.)
AMSLER	94D	PL B333 277	C. AMSler <i>et al.</i>	(Crystal Barrel Collab.)
AOYAGI	93	PL B314 246	H. Aoyagi <i>et al.</i>	(BKEI Collab.)
ARMSTRONG	93C	PL B307 394	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
BELADIDZE	93	PL B313 276	G.M. Beladidze <i>et al.</i>	(VES Collab.)
CONDO	93	PR D48 3045	G.T. Condo <i>et al.</i>	(SLAC Hybrid Collab.)
ALDE	92B	ZPHY C54 549	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP+)
BELADIDZE	92	ZPHY C54 235	G.M. Beladidze <i>et al.</i>	(VES Collab.)
ALBRECHT	90G	ZPHY C48 183	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ARMSTRONG	90	ZPHY C48 213	T.A. Armstrong, M. Benayoun, W. Beusch	(WA76 Coll.)
BARU	90	ZPHY C48 581	S.E. Baru <i>et al.</i>	(MD-1 Collab.)
BEHREND	90C	ZPHY C46 583	H.J. Behrend <i>et al.</i>	(CELLO Collab.)

BUTLER	90	PR D42 1368	F. Butler <i>et al.</i>	(Mark II Collab.)
OEST	90	ZPHY C47 343	T. Oest <i>et al.</i>	(JADE Collab.)
AUGUSTIN	89	NP B320 1	J.E. Augustin, G. Cosme	(DM2 Collab.)
VOROBYEV	88	SJNP 48 273	P.V. Vorobiev <i>et al.</i>	(NOVO)
		Translated from YAF 48 436.		
ALTHOFF	86	ZPHY C31 537	M. Althoff <i>et al.</i>	(TASSO Collab.)
ANTREASYAN	86	PR D33 1847	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
BERGER	84C	PL 149B 427	C. Berger <i>et al.</i>	(PLUTO Collab.)
BEHREND	83B	PL 125B 518 (erratum)	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
CHANGIR	82	PL 117B 123	S. Changir <i>et al.</i>	(FNAL, MINN, ROCH)
CLELAND	82B	NP B208 228	W.E. Cleland <i>et al.</i>	(DURH, GEVA, LAUS+)
EDWARDS	82F	PL 110B 82	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
DELFOSE	81	NP B183 349	A. Delfosse <i>et al.</i>	(GEVA, LAUS)
EVANGELIS...	81	NP B178 197	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
CHABAUD	80	NP B175 189	V. Chabaud <i>et al.</i>	(CERN, MPIM, AMST)
DAUM	80C	PL 89B 276	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
BALTAY	78B	PR D17 62	C. Baltay <i>et al.</i>	(COLU, BING)
CHABAUD	78	NP B145 349	V. Chabaud <i>et al.</i>	(CERN, MPIM)
FERRERSORIA	78	PL 74B 287	A. Ferrer Soria <i>et al.</i>	(ORSAY, CERN, CDF+)
HYAMS	78	NP B146 303	B.D. Hyams <i>et al.</i>	(CERN, MPIM, ATEN)
MARTIN	78D	PL 74B 417	A.D. Martin <i>et al.</i>	(DURH, GEVA, JP)
MAY	77	PR D16 1983	E.N. May <i>et al.</i>	(ROCH, CORN)
FORINO	76	NC 35A 465	A. Forino <i>et al.</i>	(BGNA, FIRZ, GENO, MILA+)
MARGULIE	76	PR D14 667	M. Margulies <i>et al.</i>	(BNL, CUNY)
EMMS	75	PL 58B 117	M.J. Emms <i>et al.</i>	(BIRM, DURH, RHIL)
WAGNER	75	PL 58B 201	F. Wagner, M. Tabak, D.M. Chew	(LBL, JP)
DIAZ	74	PRL 32 260	J. Diaz <i>et al.</i>	(CASE, CMU)
KARSHON	74	PRL 32 852	U. Karshon <i>et al.</i>	(REHO)
ANTIPOV	73	NP B63 175	Y.M. Antipov <i>et al.</i>	(CERN, SERP)
ANTIPOV	73C	NP B63 153	Y.M. Antipov <i>et al.</i>	(CERN, SERP)
CHALOUKPA	73	PL 44B 211	V. Chaloupka <i>et al.</i>	(CERN)
CONFORTO	73	PL 45B 154	G. Conforto <i>et al.</i>	(EFI, FNAL, TINTO+)
DEFOIX	73	PL 43B 141	C. Defoix <i>et al.</i>	(CDFE)
EISENSTEIN	73	PR D7 278	L. Eisenstein <i>et al.</i>	(ILL)
KEY	73	PRL 30 503	A.W. Key <i>et al.</i>	(TINTO, EFI, FNAL, WISC)
TOET	73	NP B63 248	D.Z. Toet <i>et al.</i>	(NUM, BONN, DURH, TORI)
DAMERI	72	NC 9A 1	M. Dameri <i>et al.</i>	(GENO, MILA, SACL)
EISENBERG	72	PR D5 15	Y. Eisenberg <i>et al.</i>	(REHO, SLAC, TELA)
ESPIGAT	72	NP B36 93	P. Espigat <i>et al.</i>	(CERN, CDFE)
FOLEY	72	PR D6 747	K.J. Foley <i>et al.</i>	(BNL, CUNY)
ALSTON-...	71	PL 34B 156	M. Alston-Garnjost <i>et al.</i>	(LRL)
BARNHAM	71	PRL 26 1494	K.W.J. Barnham <i>et al.</i>	(LBL)
BINNIE	71	PL 36B 257	D.M. Binnie <i>et al.</i>	(LOIC, SHMP)
BOWEN	71	PRL 26 1663	D.R. Bowen <i>et al.</i>	(NEAS, STON)
GRAY	71	PL 34B 333	G. Grayer <i>et al.</i>	(CERN, MPIM)
ABRAMOV...	70B	NP B23 466	M. Abramovich <i>et al.</i>	(CERN, JP)
ALSTON-...	70	PL 33B 607	M. Alston-Garnjost <i>et al.</i>	(LRL)
BOECKMANN	70	NP B16 221	K. Boeckmann <i>et al.</i>	(BONN, DURH, NUM+)
ASCOLI	68	PRL 20 1321	G. Ascoli <i>et al.</i>	(ILL, JP)
BOESEBECK	68	NP B4 501	K. Boesebeck <i>et al.</i>	(AACH, BERL, CERN)
CHUNG	68	PR 165 1491	S.U. Chung <i>et al.</i>	(LRL)
CONTE	67	NC 51A 175	F. Conte <i>et al.</i>	(GENO, HAMB, MILA, SACL)

 $f_0(1370)$ 

$$I^G(J^{PC}) = 0^+(0^{++})$$

See also the mini-reviews on scalar mesons under  $f_0(500)$  (see the index for the page number) and on non- $q\bar{q}$  candidates in PDG 06, Journal of Physics **G33** 1 (2006).

 $f_0(1370)$  T-MATRIX POLE POSITIONNote that  $\Gamma \approx 2 \text{Im}(\sqrt{s_{\text{pole}}})$ .

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>(1200–1500) — <math>i(150\text{--}250)</math> OUR ESTIMATE</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
(1290 ± 50) — $i(170 \pm 20)$ — 40	<sup>1</sup> ANISOVICH	09	RVUE 0.0 $\bar{p}p, \pi N$
(1373 ± 15) — $i(137 \pm 10)$	<sup>2</sup> BARGIOTTI	03	OBLX $\bar{p}p$
(1302 ± 17) — $i(166 \pm 18)$	<sup>3</sup> BARBERIS	00c	450 $\rho p \rightarrow \rho_f 4\pi p_s$
(1312 ± 25 ± 10) — $i(109 \pm 22 \pm 15)$	BARBERIS	99D	OMEG 450 $\rho p \rightarrow K^+ K^-$ , $\pi^+ \pi^-$
(1406 ± 19) — $i(80 \pm 6)$	<sup>4</sup> KAMINSKI	99	RVUE $\pi\pi \rightarrow \pi\pi, K\bar{K}, \sigma\sigma$
(1300 ± 20) — $i(120 \pm 20)$	ANISOVICH	98B	RVUE Compilation
(1290 ± 15) — $i(145 \pm 15)$	BARBERIS	97B	OMEG 450 $\rho p \rightarrow \pi\pi^0 \pi^0$
(1548 ± 40) — $i(560 \pm 40)$	BERTIN	97C	OBLX 0.0 $\bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
(1380 ± 40) — $i(180 \pm 25)$	ABELE	96B	CBAR 0.0 $\bar{p}p \rightarrow \pi^0 K_L^0 K_L^0$
(1300 ± 15) — $i(115 \pm 8)$	BUGG	96	RVUE
(1330 ± 50) — $i(150 \pm 40)$	<sup>5</sup> AMSLER	95B	CBAR $\bar{p}p \rightarrow 3\pi^0$
(1360 ± 35) — $i(150\text{--}300)$	<sup>5</sup> AMSLER	95C	CBAR $\bar{p}p \rightarrow \pi^0 \eta \eta$
(1390 ± 30) — $i(190 \pm 40)$	<sup>6</sup> AMSLER	95D	CBAR $\bar{p}p \rightarrow 3\pi^0, \pi^0 \eta \eta, \pi^0 \pi^0 \eta$
1346 — $i249$	<sup>7,8</sup> JANSSEN	95	RVUE $\pi\pi \rightarrow \pi\pi, K\bar{K}$
1214 — $i168$	<sup>8,9</sup> TORNVIST	95	RVUE $\pi\pi \rightarrow \pi\pi, K\bar{K}, K\pi, \eta\pi$
1364 — $i139$	AMSLER	94D	CBAR $\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
(1365 ± 20) — $i(134 \pm 35)$ — 55	ANISOVICH	94	CBAR $\bar{p}p \rightarrow 3\pi^0, \pi^0 \eta \eta$
(1340 ± 40) — $i(127 \pm 30)$ — 20	<sup>10</sup> BUGG	94	RVUE $\bar{p}p \rightarrow 3\pi^0, \eta \pi^0, \eta \pi^0 \pi^0$
(1430 ± 5) — $i(73 \pm 13)$	<sup>11</sup> KAMINSKI	94	RVUE $\pi\pi \rightarrow \pi\pi, K\bar{K}$
1420 — $i220$	<sup>12</sup> AU	87	RVUE $\pi\pi \rightarrow \pi\pi, K\bar{K}$

<sup>1</sup> Another pole is found at  $(1510 \pm 130) - i(800 \pm 100) \text{ MeV}$ .<sup>2</sup> Coupled channel analysis of  $\pi^+ \pi^- \pi^0, K^+ K^- \pi^0$ , and  $K^\pm K_S^0 \pi^\mp$ .<sup>3</sup> Average between  $\pi^+ \pi^- 2\pi^0$  and  $2(\pi^+ \pi^-)$ .<sup>4</sup> T-matrix pole on sheet — — —.<sup>5</sup> Supersedes ANISOVICH 94.<sup>6</sup> Coupled-channel analysis of  $\bar{p}p \rightarrow 3\pi^0, \pi^0 \eta \eta$ , and  $\pi^0 \pi^0 \eta$  on sheet IV. Demonstrates explicitly that  $f_0(500)$  and  $f_0(1370)$  are two different poles.

# Meson Particle Listings

## $f_0(1370)$

<sup>7</sup> Analysis of data from FALVARD 88.

<sup>8</sup> The pole is on Sheet III. Demonstrates explicitly that  $f_0(500)$  and  $f_0(1370)$  are two different poles.

<sup>9</sup> Uses data from BEIER 72b, OCHS 73, HYAMS 73, GRAYER 74, ROSSELET 77, CASON 83, ASTON 88, and ARMSTRONG 91b. Coupled channel analysis with flavor symmetry and all light two-pseudoscalars systems.

<sup>10</sup> Reanalysis of ANISOVICH 94 data.

<sup>11</sup> T-matrix pole on sheet III.

<sup>12</sup> Analysis of data from OCHS 73, GRAYER 74, BECKER 79, and CASON 83.

### $f_0(1370)$ BREIT-WIGNER MASS OR K-MATRIX POLE PARAMETER

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
<b>1200 to 1500 OUR ESTIMATE</b>				
<b><math>\pi\pi</math> MODE</b>				
VALUE (MeV)	EPTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1400 ± 40		<sup>1</sup> AUBERT	09L BABR	$B^\pm \rightarrow \pi^\pm \pi^\pm \pi^\mp$
1470 $^{+6}_{-7}$ $^{+72}_{-255}$		<sup>2</sup> UEHARA	08A BELL	10.6 $e^+e^- \rightarrow \pi^0 \pi^0$
1259 ± 55	2.6k	BONVICINI	07 CLEO	$D^+ \rightarrow \pi^- \pi^+ \pi^+$
1309 ± 1 ± 15		<sup>3</sup> BUGG	07A RVUE	$0.0 \ p\bar{p} \rightarrow 3\pi^0$
1449 ± 13	4.3k	<sup>4</sup> GARMASH	06 BELL	$B^+ \rightarrow K^+ \pi^+ \pi^-$
1350 ± 50		ABLIKIM	05 BES2	$J/\psi \rightarrow \phi \pi^+ \pi^-$
1265 ± 30 $^{+20}_{-35}$		ABLIKIM	05Q BES2	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^-$
1434 ± 18 ± 9	848	AITALA	01A E791	$D_S^+ \rightarrow \pi^- \pi^+ \pi^+$
1308 ± 10		BARBERIS	99B OMEG	450 $pp \rightarrow p_S p_f \pi^+ \pi^-$
1315 ± 50		BELLAZZINI	99 GAM4	450 $pp \rightarrow p p \pi^0 \pi^0$
1315 ± 30		ALDE	98 GAM4	100 $\pi^- p \rightarrow \pi^0 \pi^0 n$
1280 ± 55		BERTIN	98 OBLX	0.05-0.405 $\bar{p}p \rightarrow \pi^+ \pi^+ \pi^-$
1186		<sup>5,6</sup> TORNQVIST	95 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, K\pi, \eta\pi$
1472 ± 12		ARMSTRONG 91	OMEG	300 $pp \rightarrow p p \pi\pi, p p K\bar{K}$
1275 ± 20		BREAKSTONE 90	SFM	62 $pp \rightarrow p p \pi^+ \pi^-$
1420 ± 20		AKESSON	86 SPEC	63 $pp \rightarrow p p \pi^+ \pi^-$
1256		FROGGATT	77 RVUE	$\pi^+ \pi^-$ channel

<sup>1</sup> Breit-Wigner mass.

<sup>2</sup> Breit-Wigner mass. May also be the  $f_0(1500)$ .

<sup>3</sup> Reanalysis of ABELE 96c data.

<sup>4</sup> Also observed by GARMASH 07 in  $B^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays. Supersedes GARMASH 05.

<sup>5</sup> Uses data from BEIER 72b, OCHS 73, HYAMS 73, GRAYER 74, ROSSELET 77, CASON 83, ASTON 88, and ARMSTRONG 91b. Coupled channel analysis with flavor symmetry and all light two-pseudoscalars systems.

<sup>6</sup> Also observed by ASNER 00 in  $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$  decays

### $K\bar{K}$ MODE

VALUE (MeV)	EPTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1360 ± 31 ± 28	430	<sup>1,2</sup> DOBBS	15	$J/\psi \rightarrow \gamma K^+ K^-$
1350 ± 48 ± 15	168	<sup>1,2</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K^+ K^-$
1440 ± 6		VLADIMIRSK...06	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
1391 ± 10		TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
1440 ± 50		BOLONKIN 88	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
1463 ± 9		ETKIN 82b	MPS	23 $\pi^- p \rightarrow n 2K_S^0$
1425 ± 15		WICKLUND 80	SPEC	6 $\pi N \rightarrow K^+ K^- N$
~ 1300		POLYCHRO... 79	STRC	7 $\pi^- p \rightarrow n 2K_S^0$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.

<sup>2</sup> From a fit to a Breit-Wigner line shape with fixed  $\Gamma = 346$  MeV.

### 4 $\pi$ MODE 2( $\pi\pi$ ) $s + \rho\rho$

VALUE (MeV)	EPTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1395 ± 40		ABELE	01 CBAR	0.0 $\bar{p}d \rightarrow \pi^- 4\pi^0 p$
1374 ± 38		AMSLER	94 CBAR	0.0 $\bar{p}p \rightarrow \pi^+ \pi^- 3\pi^0$
1345 ± 12		ADAMO	93 OBLX	$\bar{\pi}p \rightarrow 3\pi^+ 2\pi^-$
1386 ± 30		GASPERO	93 DBC	0.0 $\bar{p}n \rightarrow 2\pi^+ 3\pi^-$
~ 1410	5751	<sup>1</sup> BETTINI	66 DBC	0.0 $\bar{p}n \rightarrow 2\pi^+ 3\pi^-$
<sup>1</sup> $\rho\rho$ dominant.				

### $\eta\eta$ MODE

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1262 $^{+51}_{-78}$ $^{+82}_{-103}$		<sup>1</sup> UEHARA	10A BELL	10.6 $e^+e^- \rightarrow e^+e^- \eta\eta$
1430		AMSLER	92 CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \eta\eta$
1220 ± 40		ALDE	86D GAM4	100 $\pi^- p \rightarrow n 2\eta$
<sup>1</sup> Breit-Wigner mass. May also be the $f_0(1500)$ .				

### COUPLED CHANNEL MODE

VALUE (MeV)	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1306 ± 20	<sup>1</sup> ANISOVICH 03	RVUE
<sup>1</sup> K-matrix pole from combined analysis of $\pi^- p \rightarrow \pi^0 \pi^0 n, \pi^- p \rightarrow K\bar{K}n, \pi^+ \pi^- \rightarrow \pi^+ \pi^-, \bar{p}p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta\eta, \pi^0 \pi^0 \eta, \pi^+ \pi^- \pi^0, K^+ K^- \pi^0, K_S^0 K_S^0 \pi^0, K^+ K_S^0 \pi^-$ at rest, $\bar{p}n \rightarrow \pi^- \pi^- \pi^+, K_S^0 K^- \pi^0, K_S^0 K_S^0 \pi^-$ at rest.		

### $f_0(1370)$ BREIT-WIGNER WIDTH

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
<b>200 to 500 OUR ESTIMATE</b>				
<b><math>\pi\pi</math> MODE</b>				
VALUE (MeV)	EPTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
300 ± 80		<sup>1</sup> AUBERT	09L BABR	$B^\pm \rightarrow \pi^\pm \pi^\pm \pi^\mp$
90 $^{+2}_{-1}$ $^{+50}_{-22}$		<sup>2</sup> UEHARA	08A BELL	10.6 $e^+e^- \rightarrow e^+e^- \pi^0 \pi^0$
298 ± 21	2.6k	BONVICINI	07 CLEO	$D^+ \rightarrow \pi^- \pi^+ \pi^+$
126 ± 25	4286	<sup>3</sup> GARMASH	06 BELL	$B^+ \rightarrow K^+ \pi^+ \pi^-$
265 ± 40		ABLIKIM	05 BES2	$J/\psi \rightarrow \phi \pi^+ \pi^-$
350 ± 100 $^{+105}_{-60}$		ABLIKIM	05Q BES2	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^-$
173 ± 32 ± 6	848	AITALA	01A E791	$D_S^+ \rightarrow \pi^- \pi^+ \pi^+$
222 ± 20		BARBERIS	99B OMEG	450 $pp \rightarrow p_S p_f \pi^+ \pi^-$
255 ± 60		BELLAZZINI	99 GAM4	450 $pp \rightarrow p p \pi^0 \pi^0$
190 ± 50		ALDE	98 GAM4	100 $\pi^- p \rightarrow \pi^0 \pi^0 n$
323 ± 13		BERTIN	98 OBLX	0.05-0.405 $\bar{p}p \rightarrow \pi^+ \pi^+ \pi^-$
350		<sup>4,5</sup> TORNQVIST	95 RVUE	$\pi\pi \rightarrow \pi\pi, K\bar{K}, K\pi, \eta\pi$
195 ± 33		ARMSTRONG 91	OMEG	300 $pp \rightarrow p p \pi\pi, p p K\bar{K}$
285 ± 60		BREAKSTONE 90	SFM	62 $pp \rightarrow p p \pi^+ \pi^-$
460 ± 50		AKESSON	86 SPEC	63 $pp \rightarrow p p \pi^+ \pi^-$
~ 400		<sup>6</sup> FROGGATT	77 RVUE	$\pi^+ \pi^-$ channel

<sup>1</sup> The systematic errors are not reported.

<sup>2</sup> Breit-Wigner width. May also be the  $f_0(1500)$ .

<sup>3</sup> Also observed by GARMASH 07 in  $B^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays. Supersedes GARMASH 05.

<sup>4</sup> Uses data from BEIER 72b, OCHS 73, HYAMS 73, GRAYER 74, ROSSELET 77, CASON 83, ASTON 88, and ARMSTRONG 91b. Coupled channel analysis with flavor symmetry and all light two-pseudoscalars systems.

<sup>5</sup> Also observed by ASNER 00 in  $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$  decays

<sup>6</sup> Width defined as distance between 45 and 135° phase shift.

### $K\bar{K}$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
121 ± 15	VLADIMIRSK...06	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
55 ± 26	TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
250 ± 80	BOLONKIN 88	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
118 $^{+138}_{-16}$	ETKIN 82b	MPS	23 $\pi^- p \rightarrow n 2K_S^0$
160 ± 30	WICKLUND 80	SPEC	6 $\pi N \rightarrow K^+ K^- N$
~ 150	POLYCHRO... 79	STRC	7 $\pi^- p \rightarrow n 2K_S^0$

### 4 $\pi$ MODE 2( $\pi\pi$ ) $s + \rho\rho$

VALUE (MeV)	EPTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
275 ± 55		ABELE	01 CBAR	0.0 $\bar{p}d \rightarrow \pi^- 4\pi^0 p$
375 ± 61		AMSLER	94 CBAR	0.0 $\bar{p}p \rightarrow \pi^+ \pi^- 3\pi^0$
398 ± 26		ADAMO	93 OBLX	$\bar{\pi}p \rightarrow 3\pi^+ 2\pi^-$
310 ± 50		GASPERO	93 DBC	0.0 $\bar{p}n \rightarrow 2\pi^+ 3\pi^-$
~ 90	5751	<sup>1</sup> BETTINI	66 DBC	0.0 $\bar{p}n \rightarrow 2\pi^+ 3\pi^-$
<sup>1</sup> $\rho\rho$ dominant.				

### $\eta\eta$ MODE

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
484 $^{+246}_{-170}$ $^{+246}_{-263}$		<sup>1</sup> UEHARA	10A BELL	10.6 $e^+e^- \rightarrow e^+e^- \eta\eta$
250		AMSLER	92 CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \eta\eta$
320 ± 40		ALDE	86D GAM4	100 $\pi^- p \rightarrow n 2\eta$
<sup>1</sup> Breit-Wigner width. May also be the $f_0(1500)$ .				

### COUPLED CHANNEL MODE

VALUE (MeV)	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •		
147 $^{+30}_{-50}$	<sup>1</sup> ANISOVICH 03	RVUE
<sup>1</sup> K-matrix pole from combined analysis of $\pi^- p \rightarrow \pi^0 \pi^0 n, \pi^- p \rightarrow K\bar{K}n, \pi^+ \pi^- \rightarrow \pi^+ \pi^-, \bar{p}p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta\eta, \pi^0 \pi^0 \eta, \pi^+ \pi^- \pi^0, K^+ K^- \pi^0, K_S^0 K_S^0 \pi^0, K^+ K_S^0 \pi^-$ at rest, $\bar{p}n \rightarrow \pi^- \pi^- \pi^+, K_S^0 K^- \pi^0, K_S^0 K_S^0 \pi^-$ at rest.		



$f_0(1370)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi\pi$	seen
$\Gamma_2$ $4\pi$	seen
$\Gamma_3$ $4\pi^0$	seen
$\Gamma_4$ $2\pi^+2\pi^-$	seen
$\Gamma_5$ $\pi^+\pi^-2\pi^0$	seen
$\Gamma_6$ $\rho\rho$	dominant
$\Gamma_7$ $2(\pi\pi)$ s-wave	seen
$\Gamma_8$ $\pi(1300)\pi$	seen
$\Gamma_9$ $a_1(1260)\pi$	seen
$\Gamma_{10}$ $\eta\eta$	seen
$\Gamma_{11}$ $K\bar{K}$	seen
$\Gamma_{12}$ $K\bar{K}n\pi$	not seen
$\Gamma_{13}$ $6\pi$	not seen
$\Gamma_{14}$ $\omega\omega$	not seen
$\Gamma_{15}$ $\gamma\gamma$	seen
$\Gamma_{16}$ $e^+e^-$	not seen

 $f_0(1370)$  PARTIAL WIDTHS

$\Gamma(\gamma\gamma)$	See $\gamma\gamma$ widths under $f_0(500)$ and MORGAN 90.				$\Gamma_{15}$
$\Gamma(e^+e^-)$					$\Gamma_{16}$
<u>VALUE (eV)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;20</b>	90	VOROBYEV	88	ND $e^+e^- \rightarrow \pi^0\pi^0$	

 $f_0(1370)$   $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

$\Gamma(\eta\eta) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_{10}\Gamma_{15}/\Gamma$		
VALUE (eV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$121^{+133+169}_{-53-106}$	<sup>1</sup> UEHARA	10A	BELL $10.6 e^+e^- \rightarrow e^+e^-\eta\eta$
<sup>1</sup> Including interference with the $f_2'(1525)$ (parameters fixed to the values from the 2008 edition of this review, PDG 08) and $f_2(1270)$ . May also be the $f_0(1500)$ .			

 $f_0(1370)$  BRANCHING RATIOS

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.26 \pm 0.09$	BUGG	96	RVUE
$<0.15$	<sup>1</sup> AMSLER	94	CBAR $\bar{p}p \rightarrow \pi^+ \pi^- 3\pi^0$
$<0.06$	GASPERO	93	DBC $0.0 \bar{p}n \rightarrow \text{hadrons}$
<sup>1</sup> Using AMSLER 95B ( $3\pi^0$ ).			

$\Gamma(4\pi)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma = (\Gamma_3+\Gamma_4+\Gamma_5)/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
>0.72	GASPERO	93	DBC 0.0 $\bar{p}n \rightarrow$ hadrons

$\Gamma(4\pi^0)/\Gamma(4\pi)$				$\Gamma_3/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	ABELE	96	CBAR	$0.0 \bar{p}p \rightarrow 5\pi^0$
$0.068 \pm 0.005$	<sup>1</sup> GASPERO	93	DBC	$0.0 \bar{p}n \rightarrow \text{hadrons}$
<sup>1</sup> Model-dependent evaluation.				

$\Gamma(2\pi^+2\pi^-)/\Gamma(4\pi)$	$\Gamma_4/\Gamma_2 = \Gamma_4/(\Gamma_3+\Gamma_4+\Gamma_5)$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.420 \pm 0.014$	<sup>1</sup> GASPERO	93	DBC $0.0 \bar{p}n \rightarrow 2\pi^+3\pi^-$
<sup>1</sup> Model-dependent evaluation.			

$\Gamma(\pi^+\pi^-2\pi^0)/\Gamma(4\pi)$	$\Gamma_5/\Gamma_2=\Gamma_5/(\Gamma_3+\Gamma_4+\Gamma_5)$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.512±0.019	<sup>1</sup> GASPERO	93	DBC 0.0 $\bar{p}n \rightarrow$ hadrons
<sup>1</sup> Model-dependent evaluation.			

$\Gamma(\rho\rho)/\Gamma(4\pi)$				$\Gamma_6/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.26 \pm 0.07$	ABELE	01B	CBAR $0.0 \bar{p}d \rightarrow 5\pi p$	

 $\Gamma(2(\pi\pi)\text{s-wave})/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$5.6 \pm 2.6$	<sup>1</sup> ABELE	01	CBAR $0.0 \bar{p}d \rightarrow \pi^-4\pi^0p$
<sup>1</sup> From the combined data of ABLE 96 and ABLE 96c.			

 $\Gamma(2(\pi\pi)\text{s-wave})/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.51 \pm 0.09$	ABELE	01B	CBAR $0.0 \bar{p}d \rightarrow 5\pi p$

 $\Gamma(\rho\rho)/\Gamma(2(\pi\pi)\text{s-wave})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
large	BARBERIS	00C	$450 p p \rightarrow p_f 4\pi p_s$
$1.6 \pm 0.2$	AMSLER	94	CBAR $\bar{p}p \rightarrow \pi^+\pi^-3\pi^0$
$\sim 0.65$	GASPERO	93	DBC $0.0 \bar{p}n \rightarrow \text{hadrons}$

 $\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.17 \pm 0.06$	ABELE	01B	CBAR $0.0 \bar{p}d \rightarrow 5\pi p$

 $\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.06 \pm 0.02$	ABELE	01B	CBAR $0.0 \bar{p}d \rightarrow 5\pi p$

 $\Gamma(\eta\eta)/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$(28 \pm 11) \times 10^{-3}$	<sup>1</sup> ANISOVICH	02D	SPEC Combined fit
$(4.7 \pm 2.0) \times 10^{-3}$	BARBERIS	00E	$450 p p \rightarrow p_f \eta \eta p_s$
<sup>1</sup> From a combined K-matrix analysis of Crystal Barrel ( $0. p \bar{p} \rightarrow \pi^0\pi^0\pi^0, \pi^0\eta\eta, \pi^0\pi^0\eta$ ), GAMS ( $\pi p \rightarrow \pi^0\pi^0n, \eta\eta n, \eta\eta' n$ ), and BNL ( $\pi p \rightarrow K\bar{K}n$ ) data.			

 $\Gamma(K\bar{K})/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.35 \pm 0.13$	BUGG	96	RVUE

 $\Gamma(K\bar{K}n\pi)/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.08 \pm 0.08$	ABLIKIM	05	BES2 $J/\psi \rightarrow \phi\pi^+\pi^-, \phi K^+K^-$
$0.91 \pm 0.20$	<sup>1</sup> BARGIOTTI	03	OBLX $\bar{p}p$
$0.12 \pm 0.06$	<sup>2</sup> ANISOVICH	02D	SPEC Combined fit
$0.46 \pm 0.15 \pm 0.11$	BARBERIS	99D	OMEG $450 p p \rightarrow K^+K^-, \pi^+\pi^-$
<sup>1</sup> Coupled channel analysis of $\pi^+\pi^-\pi^0, K^+K^-\pi^0$ , and $K^\pm K_S^0 \pi^\mp$ .			
<sup>2</sup> From a combined K-matrix analysis of Crystal Barrel ( $0. p \bar{p} \rightarrow \pi^0\pi^0\pi^0, \pi^0\eta\eta, \pi^0\pi^0\eta$ ), GAMS ( $\pi p \rightarrow \pi^0\pi^0n, \eta\eta n, \eta\eta' n$ ), and BNL ( $\pi p \rightarrow K\bar{K}n$ ) data.			

 $\Gamma(K\bar{K}n\pi)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$<0.03$	GASPERO	93	DBC $0.0 \bar{p}n \rightarrow \text{hadrons}$

 $\Gamma(6\pi)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$<0.22$	GASPERO	93	DBC $0.0 \bar{p}n \rightarrow \text{hadrons}$

 $\Gamma(\omega\omega)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$<0.13$	GASPERO	93	DBC $0.0 \bar{p}n \rightarrow \text{hadrons}$

 $f_0(1370)$  REFERENCES

DOBBS	15	PR D91 052006	S. Dobbs <i>et al.</i>	(NWES)
UEHARA	10A	PR D82 114031	S. Uehara <i>et al.</i>	(BELLE Collab.)
ANISOVICH	09	IJMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
AUBERT	09L	PR D79 072006	B. Aubert <i>et al.</i>	(BABAR Collab.)
PDG	08	PL B667 1	C. Amisler <i>et al.</i>	(PDG Collab.)
UEHARA	08A	PR D78 052004	S. Uehara <i>et al.</i>	(BELLE Collab.)
BONVICINI	07	PR D76 012001	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
BUGG	07A	JP G34 151	D.V. Bugg <i>et al.</i>	
GARMASH	07	PR D75 012006	A. Garmash <i>et al.</i>	(BELLE Collab.)
GARMASH	06	PRL 96 251803	A. Garmash <i>et al.</i>	(BELLE Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
VLADIMIRSK...	06	PAN 69 493	V.V. Vladimirov <i>et al.</i>	(ITEP, Moscow)
Translated from YAF 69 515.				

Meson Particle Listings

$f_0(1370)$ ,  $h_1(1380)$ ,  $\pi_1(1400)$

ABLIKIM	05	PL B607 243	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05Q	PR D72 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
GARMASH	05	PR D71 092003	A. Garmash <i>et al.</i>	(BELLE Collab.)
ANISOVICH	03	EPJ A16 229	V.V. Anisovich <i>et al.</i>	
BARGIOTTI	03	EPJ C26 371	M. Bargiotti <i>et al.</i>	(OBELIX Collab.)
TIKHOMIROV	03	PAN 66 828	G.D. Tikhomirov <i>et al.</i>	
ANISOVICH	02D	Translated from YAF 66 860	V.V. Anisovich <i>et al.</i>	
ABELE	01	EPJ C19 667	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	01B	EPJ C21 261	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
AITALA	01A	PRL 86 765	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
ASNER	00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BARBERIS	00C	PL B471 440	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00E	PL B479 59	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	99B	PL B453 316	D. Barberis <i>et al.</i>	(Omega Expt.)
BARBERIS	99D	PL B462 462	D. Barberis <i>et al.</i>	(Omega Expt.)
BELLAZZINI	99	PL B467 296	R. Bellazzini <i>et al.</i>	
KAMINSKI	99	EPJ C9 141	R. Kaminski, L. Lesniak, B. Loiseau	(CRAC, PARIN)
ALDE	98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
Also		PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
ANISOVICH	98B	Translated from YAF 62 446	V.V. Anisovich <i>et al.</i>	
BERTIN	98	SFU 41 419	V.V. Anisovich <i>et al.</i>	
BARBERIS	97B	Translated from UFN 168 481	A. Bertin <i>et al.</i>	(OBELIX Collab.)
BERTIN	97C	PR D57 55	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ABELE	96	PL B380 453	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	96B	PL B385 425	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	96C	NP A609 562	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BUGG	96	NP B471 59	D.V. Bugg, A.V. Sarantsev, B.S. Zou	(LOQM, PNPI)
AMSLER	95B	PL B342 433	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	95C	PL B353 571	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	95D	PL B355 425	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
JANSEN	95	PR D52 2690	G. Jansen <i>et al.</i>	(STON, ADLD, JULI)
TORNQVIST	95	ZPHY C68 647	N.A. Tornqvist	(HELS)
AMSLER	94	PL B322 431	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	94D	PL B333 277	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
ANISOVICH	94	PL B323 233	V.V. Anisovich <i>et al.</i>	(Crystal Barrel Collab.)
BUGG	94	PR D50 4412	D.V. Bugg <i>et al.</i>	(LOQM)
KAMINSKI	94	PR D50 3145	R. Kaminski, L. Lesniak, J.P. Maillet	(CRAC+)
ADAMO	93	NP A558 13C	A. Adamo <i>et al.</i>	(OBELIX Collab.)
GASPERO	93	NP A562 407	M. Gaspero	(ROMA1)JPC
AMSLER	92	PL B291 347	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
ARMSTRONG	91	ZPHY C51 351	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
ARMSTRONG	91B	ZPHY C52 389	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
BREAKSTONE	90	ZPHY C48 569	A.M. Breakstone <i>et al.</i>	(ISU, BGNA, CERN+)
MORGAN	90	ZPHY C48 623	D. Morgan, M.R. Pennington	(RAL, DURH)
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
BOLONKIN	88	NP B309 426	B.V. Bolonkin <i>et al.</i>	(ITEP, SERP)
FALVARD	88	PR D38 2706	A. Falvard <i>et al.</i>	(CLER, FRAS, LALO+)
VOROBYEV	88	SJNP 48 273	P.V. Vorobiev <i>et al.</i>	(NOVO)
AU	87	PR D35 1633	K.L. Au, D. Morgan, M.R. Pennington	(DURH, RAL)
AKESSON	86	NP B264 154	T. Akesson <i>et al.</i>	(Axial Field Spec. Collab.)
ALDE	86D	NP B269 405	D.M. Alde <i>et al.</i>	(BELG, LAPP, SERP, CERN+)
CAS ON	83	PR D28 1586	N.M. Cason <i>et al.</i>	(NDAM, ANL)
ETKIN	82B	PR D25 1786	A. Etkin <i>et al.</i>	(BNL, CUNY, TUFTS, VAND)
WICKLUND	80	PRL 45 1469	A.B. Wicklund <i>et al.</i>	(ANL)
BECKER	79	NP B151 46	H. Becker <i>et al.</i>	(MPIM, CERN, ZEEM, CRAC)
POLYCHRO...	79	PR D19 1317	V.A. Polychronakos <i>et al.</i>	(NDAM, ANL)
FROGGATT	77	NP B129 89	C.D. Froggatt, J.L. Petersen	(GLAS, NORD)
ROSSELET	77	PR D15 574	L. Rosselet <i>et al.</i>	(GEVA, SACL)
GRAYER	74	NP B75 189	G. Grayer <i>et al.</i>	(CERN, MPIM)
HYAMS	73	NP B64 134	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
OCHS	73	Thesis	W. Ochs	(MPIM, MUNI)
BEIER	72B	PRL 29 511	E.W. Beier <i>et al.</i>	(PENN)
BETTINI	66	NC 42A 635	A. Bettini <i>et al.</i>	(PADO, PISA)

$h_1(1380)$

$I^G(J^{PC}) = ?^-(1^+ -)$

OMITTED FROM SUMMARY TABLE  
Seen in partial-wave analysis of the  $K\bar{K}\pi$  system. Needs confirmation.

$h_1(1380)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1407±12 OUR AVERAGE</b>	Error includes scale factor of 1.5.		
1412± 4±8	ABLIKIM	15M BES3	$\psi(2S) \rightarrow \gamma \chi_{c1,2} \rightarrow \gamma K^* \bar{K}$
1440±60	ABELE	97H CBAR	$\bar{p}p \rightarrow K_L^0 K_S^0 \pi^0 \pi^0$
1380±20	ASTON	88c LASS	$11 K^- p \rightarrow K_S^0 K^\pm \pi^\mp \Lambda$

$h_1(1380)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>89±23 OUR AVERAGE</b>			
84±12±40	ABLIKIM	15M BES3	$\psi(2S) \rightarrow \gamma \chi_{c1,2} \rightarrow \gamma K^* \bar{K}$
170±80	ABELE	97H CBAR	$\bar{p}p \rightarrow K_L^0 K_S^0 \pi^0 \pi^0$
80±30	ASTON	88c LASS	$11 K^- p \rightarrow K_S^0 K^\pm \pi^\mp \Lambda$

$h_1(1380)$  DECAY MODES

Mode	
$\Gamma_1$	$K\bar{K}^*(892) + c.c.$

$h_1(1380)$  REFERENCES

ABLIKIM	15M	PR D91 112008	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABELE	97H	PL B415 280	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ASTON	88C	PL B201 573	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)

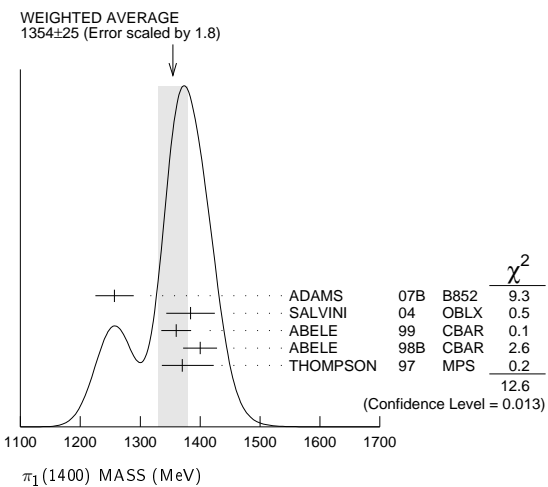
$\pi_1(1400)$

$I^G(J^{PC}) = 1^-(1^- +)$

See also the mini-review under non- $q\bar{q}$  candidates in PDG 06, Journal of Physics **G33** 1 (2006).

$\pi_1(1400)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1354 ±25 OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.				
1257 ±20 ±25	23.5k	ADAMS	07B	B852	$18 \pi^- p \rightarrow \eta \pi^0 n$
1384 ±20 ±35	90k	SALVINI	04	OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
1360 ±25		ABELE	99	CBAR	$0.0 \bar{p}p \rightarrow \pi^0 \pi^0 \eta$
1400 ±20 ±20		ABELE	98B	CBAR	$0.0 \bar{p}n \rightarrow \pi^- \pi^0 \eta$
1370 ±16 ±30		<sup>1</sup> THOMPSON	97	MPS	$18 \pi^- p \rightarrow \eta \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1323.1± 4.6		<sup>2</sup> AOYAGI	93	BKEI	$\pi^- p \rightarrow \eta \pi^- p$
1406 ±20		<sup>3</sup> ALDE	88B	GAM4	0 $100 \pi^- p \rightarrow \eta \pi^0 n$
<sup>1</sup> Natural parity exchange, questioned by DZIERBA 03. <sup>2</sup> Unnatural parity exchange. <sup>3</sup> Seen in the $P_0$ -wave intensity of the $\eta \pi^0$ system, unnatural parity exchange.					



$\pi_1(1400)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>330 ±35 OUR AVERAGE</b>					
354 ±64 ±58	23.5k	ADAMS	07B	B852	$18 \pi^- p \rightarrow \eta \pi^0 n$
378 ±50 ±50	90k	SALVINI	04	OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
220 ±90		ABELE	99	CBAR	$0.0 \bar{p}p \rightarrow \pi^0 \pi^0 \eta$
310 ±50 ±50		ABELE	98B	CBAR	$0.0 \bar{p}n \rightarrow \pi^- \pi^0 \eta$
385 ±40 ±65		<sup>4</sup> THOMPSON	97	MPS	$18 \pi^- p \rightarrow \eta \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
143.2±12.5		<sup>5</sup> AOYAGI	93	BKEI	$\pi^- p \rightarrow \eta \pi^- p$
180 ±20		<sup>6</sup> ALDE	88B	GAM4	0 $100 \pi^- p \rightarrow \eta \pi^0 n$
<sup>4</sup> Resolution is not unfolded, natural parity exchange, questioned by DZIERBA 03. <sup>5</sup> Unnatural parity exchange. <sup>6</sup> Seen in the $P_0$ -wave intensity of the $\eta \pi^0$ system, unnatural parity exchange.					

$\pi_1(1400)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$\eta \pi^0$ seen
$\Gamma_2$	$\eta \pi^-$ seen
$\Gamma_3$	$\eta' \pi$

$\pi_1(1400)$  BRANCHING RATIOS

$\Gamma(\eta \pi^0)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_1/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
not seen		PROKOSHKIN	95B	GAM4		$100 \pi^- p \rightarrow \eta \pi^0 n$
not seen		<sup>7</sup> BUGG	94	RVUE		$\bar{p}p \rightarrow \eta 2\pi^0$
not seen		<sup>8</sup> APEL	81	NICE	0	$40 \pi^- p \rightarrow \eta \pi^0 n$

<sup>7</sup> Using Crystal Barrel data.  
<sup>8</sup> A general fit allowing  $S$ ,  $D$ , and  $P$  waves (including  $m=0$ ) is not done because of limited statistics.

$\Gamma(\eta\pi^-)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •							
possibly seen	BELADIDZE	93	VES	$37\pi^- N \rightarrow \eta\pi^- N$			
$\Gamma(\eta'\pi)/\Gamma(\eta\pi^0)$				$\Gamma_3/\Gamma_1$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	VALUE	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •							
<0.80	95	BOUTEMEUR	90	GAM4	$100\pi^- p \rightarrow 4\gamma n$		

 **$\pi_1(1400)$  REFERENCES**

ADAMS	07B	PL B657 27	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
SALVINI	04	EPJ C35 21	P. Salvini <i>et al.</i>	(OBELIX Collab.)
DZIERBA	03	PR D67 094015	A.R. Dzierba <i>et al.</i>	
ABELE	99	PL B446 349	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	98B	PL B423 175	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
THOMPSON	97	PRL 79 1630	D.R. Thompson <i>et al.</i>	(BNL E852 Collab.)
PROKOSHIN	95B	PAN 58 606	Y.D. Prokoshin, S.A. Sadovsky	(SERP)
		Translated from YAF 58 662		
BUGG	94	PR D50 4412	D.V. Bugg <i>et al.</i>	(LOQM)
AOYAGI	93	PL B314 246	H. Aoyagi <i>et al.</i>	(BKEI Collab.)
BELADIDZE	93	PL B313 276	G.M. Beladidze <i>et al.</i>	(VES Collab.)
BOUTEMEUR	90	Hadron 89 Conf. p 119	M. Boutemeur, M. Poulet	(SERP, BELG, LANL, LAPP)
ALDE	88B	PL B205 397	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP)
APEL	81	NP B193 269	W.D. Apel <i>et al.</i>	(SERP, CERN)

 **$\eta(1405)$** 

$$J^G(J^{PC}) = 0^+(0^{-+})$$

See also the  $\eta(1475)$ .

**THE PSEUDOSCALAR AND PSEUDOVECTOR MESONS IN THE 1400 MEV REGION**

Revised July 2015 by C. Amsler (University of Bern) and A. Masoni (INFN Cagliari).

This minireview deals with some of the  $0^{-+}$  and  $1^{++}$  mesons reported in the 1200–1500 MeV region, namely the  $\eta(1405)$ ,  $\eta(1475)$ ,  $f_1(1285)$ ,  $f_1(1420)$ ,  $a_1(1420)$  and  $f_1(1510)$ . The first observation of a pseudoscalar resonance around 1400 MeV – the  $\eta(1440)$  – was made in  $p\bar{p}$  annihilation at rest into  $\eta(1440)\pi^+\pi^-$ ,  $\eta(1440) \rightarrow K\bar{K}\pi$  [1]. This state was reported to decay into  $a_0(980)\pi$  and  $K^*(892)\bar{K}$  with roughly equal contributions. The  $\eta(1440)$  was also observed in radiative  $J/\psi(1S)$  decay into  $K\bar{K}\pi$  [2–4] and  $\gamma\rho$  [5]. However, two pseudoscalars are now reported in this mass region, the  $\eta(1405)$  and  $\eta(1475)$ . The former decays mainly through  $a_0(980)\pi$  (or direct  $K\bar{K}\pi$ ) and the latter mainly to  $K^*(892)\bar{K}$ .

The simultaneous observation of two pseudoscalars is reported in three production mechanisms:  $\pi^-p$  [6,7]; radiative  $J/\psi(1S)$  decay [8,9]; and  $\bar{p}p$  annihilation at rest [10–13]. All of them give values for the masses, widths, and decay modes that are in reasonable agreement. However, Ref. [9] favors a state decaying into  $K^*(892)\bar{K}$  at a lower mass than the state decaying into  $a_0(980)\pi$ . In  $J/\psi(1S)$  radiative decay, the  $\eta(1405)$  decays into  $K\bar{K}\pi$  through  $a_0(980)\pi$ , and hence a signal is also expected in the  $\eta\pi\pi$  mass spectrum. This was indeed observed by MARK III in  $\eta\pi^+\pi^-$  [14], which reported a mass of 1400 MeV, in line with the existence of the  $\eta(1405)$  decaying into  $a_0(980)\pi$ .

BESII [15] observes an enhancement in  $K^+K^-\pi^0$  around 1.44 GeV in  $J/\psi(1S)$  decay, recoiling against an  $\omega$  (but not a  $\phi$ ) without resolving the presence of two states nor performing a spin-parity analysis, due to low statistics. This state could also be the  $f_1(1420)$  (see below). On the other hand, BESII observes  $\eta(1405) \rightarrow \eta\pi\pi$  in  $J/\psi(1S)$  decay, recoiling against an  $\omega$  [16]. A single unresolved broad peak is also observed by

BESIII in the decay  $\psi(2S) \rightarrow \omega K^* K$  which could be due to  $\eta(1405)$ ,  $\eta(1475)$  and  $f_1(1420)$  [17].

The  $\eta(1405)$  is also observed in  $\bar{p}p$  annihilation at rest into  $\eta\pi^+\pi^-\pi^0\pi^0$ , where it decays into  $\eta\pi\pi$  [18]. The intermediate  $a_0(980)\pi$  accounts for roughly half of the  $\eta\pi\pi$  signal, in agreement with MARK III [14] and DM2 [4].

However, the issue remains controversial as to whether two pseudoscalar mesons really exist. According to Ref. [19] the splitting of a single state could be due to nodes in the decay amplitudes which differ in  $\eta\pi\pi$  and  $K^*(892)\bar{K}$ . Based on the isospin violating decay  $J/\psi(1S) \rightarrow \gamma 3\pi$  observed by BESIII [20] the splitting could also be due to a triangular singularity mixing  $\eta\pi\pi$  and  $K^*(892)\bar{K}$  [21–22]. However, in a further paper [23], using the approach of [21], the authors concluded that the BESIII results can be reproduced either with the  $\eta(1405)$  or the  $\eta(1475)$ , or by a mixture of these two states.

The  $\eta(1295)$  has been observed by four  $\pi^-p$  experiments [7,24–26], and evidence is reported in  $\bar{p}p$  annihilation [27–29]. In  $J/\psi(1S)$  radiative decay, the  $\eta(1295)$  signal is evident in the  $0^{-+}$   $\eta\pi\pi$  wave of the DM2 data [9]. Also BaBar [30] reports evidence for a signal around 1295 MeV in  $B$  decays into  $\eta\pi\pi K$ . Nonetheless, the existence of the  $\eta(1295)$  is questioned in Refs. [19] and [31] in which the authors claim the existence of a single pseudoscalar meson at 1440 MeV, the first radial excitation of the  $\eta$ . This conclusion is mainly based on the analysis of the annihilation  $\bar{p}p \rightarrow 4\pi\eta$  with Crystal Barrel data [32].

Considering that the  $\eta(1295)$  has been reported by several experiments, using different production mechanisms, we shall assume that this state is established. The  $\eta(1475)$  could then be the first radial excitation of the  $\eta'$ , with the  $\eta(1295)$  being the first radial excitation of the  $\eta$ . Ideal mixing, suggested by the  $\eta(1295)$  and  $\pi(1300)$  mass degeneracy, would then imply that the second isoscalar in the nonet is mainly  $s\bar{s}$ , and hence couples to  $K^*\bar{K}$ , in agreement with properties of the  $\eta(1475)$ . Also, its width matches the expected width for the radially excited  $s\bar{s}$  state [33,34]. A study of radial excitations of pseudoscalar mesons [35] favors the  $s\bar{s}$  interpretation of the  $\eta(1475)$ . However, due to the strong kinematical suppression the data are not sufficient to exclude a sizeable  $s\bar{s}$  admixture also in the  $\eta(1405)$ .

The  $K\bar{K}\pi$  and  $\eta\pi\pi$  channels were studied in  $\gamma\gamma$  collisions by L3 [36]. The analysis led to a clear  $\eta(1475)$  signal in  $K\bar{K}\pi$ , decaying into  $K^*\bar{K}$ , very well identified in the untagged data sample, where contamination from spin 1 resonances is not allowed. At the same time, L3 [36] did not observe the  $\eta(1405)$ , neither in  $K\bar{K}\pi$  nor in  $\eta\pi\pi$ . The observation of the  $\eta(1475)$ , combined with the absence of an  $\eta(1405)$  signal, strengthens the two-resonances hypothesis. Since gluonium production is presumably suppressed in  $\gamma\gamma$  collisions, the L3 results [36] suggest that  $\eta(1405)$  has a large gluonic content (see also Refs. [37] and [38]).

The L3 result is somewhat in disagreement with that of CLEO-II, which did not observe any pseudoscalar signal in  $\gamma\gamma \rightarrow \eta(1475) \rightarrow K_S^0 K^\pm \pi^\mp$  [39]. However, more data are

# Meson Particle Listings

## $\eta(1405)$

required. Moreover, after the CLEO-II result, L3 performed a further analysis with full statistics [40], confirming their previous evidence for the  $\eta(1475)$ . The CLEO upper limit [39] for  $\Gamma_{\gamma\gamma}(\eta(1475))$ , and the L3 results [40], are consistent with the world average for the  $\eta(1475)$  width.

BaBar [30] also reports the  $\eta(1475)$  in  $B$  decays into  $K\bar{K}^*$  recoiling against a  $K$ , but upper limits only are given for the  $\eta(1405)$ . As mentioned above, in  $B$  decays into  $\eta\pi\pi K$  the  $\eta(1295) \rightarrow \eta\pi\pi$  is observed while only upper limits are given for the  $\eta(1405)$ . The  $f_1(1420)$  (and  $f_1(1285)$ ) are not seen.

The gluonium interpretation for the  $\eta(1405)$  is not favored by lattice gauge theories which predict the  $0^{-+}$  state above 2 GeV [41,42] (see also the article on the “Quark model” in this issue of the Review). However, the  $\eta(1405)$  is an excellent candidate for the  $0^{-+}$  glueball in the fluxtube model [43]. In this model, the  $0^{++}$   $f_0(1500)$  glueball is also naturally related to a  $0^{-+}$  glueball with mass degeneracy broken in QCD. Also, Ref. [44] shows that the pseudoscalar glueball could lie at a lower mass than predicted from lattice calculation. In this model the  $\eta(1405)$  appears as the natural glueball candidate, see also Refs. [45–47]. A detailed review of the experimental situation is available in Ref. [48].

Let us now deal with the  $1^{++}$  mesons. The pseudovector nonet is believed to consist of the isovector  $a_1(1260)$ , the isoscalars  $f_1(1285)$  and  $f_1(1420)$ , and the  $K_{1A}$ , which is a mixture of about 50%  $K_1(1270)$  and 50%  $K_1(1400)$ . (This last property prevents a straightforward calculation of the nonet mixing angle via the mass formulae.) The  $f_1(1285)$  could also be a  $K^*\bar{K}$  molecule [49] or as a tetraquark state [50] and the  $f_1(1420)$  a  $K^*\bar{K}$  molecule, due to the proximity of the  $K^*\bar{K}$  threshold [51]. LHCb has analyzed the decays  $\bar{B}^0$  and  $\bar{B}_s^0 \rightarrow J/\psi(1S)f_1(1285)$  and determined the nonet mixing angle to be consistent with a mostly  $u\bar{u} + d\bar{d}$  structure [52] without specifying the identity of its isoscalar partner. This is consistent with earlier determinations assuming the  $f_1(1420)$  as the isoscalar partner [53] and the ratio of  $\bar{B}^0/\bar{B}_s^0$  decay rates excludes the tetraquark interpretation of this state [52].

The  $f_1(1420)$ , decaying into  $K^*\bar{K}$ , was first reported in  $\pi^-p$  reactions at 4 GeV/c [54]. However, later analyses found that the 1400–1500 MeV region was far more complex [55–57]. A reanalysis of the MARK III data in radiative  $J/\psi(1S)$  decay into  $K\bar{K}\pi$  [8] shows the  $f_1(1420)$  decaying into  $K^*\bar{K}$ . A  $C=+1$  state is also seen in tagged  $\gamma\gamma$  collisions (*e.g.*, Ref. [58]).

In  $\pi^-p \rightarrow \eta\pi\pi n$  charge-exchange reactions at 8–9 GeV/c the  $\eta\pi\pi$  mass spectrum is dominated by the  $\eta(1440)$  and  $\eta(1295)$  [24,59], and at 100 GeV/c Ref. [25] reports the  $\eta(1295)$  and  $\eta(1440)$  decaying into  $\eta\pi^0\pi^0$  with a weak  $f_1(1285)$  signal, and no evidence for the  $f_1(1420)$ .

Axial ( $1^{++}$ ) mesons are not observed in  $\bar{p}p$  annihilation at rest in liquid hydrogen, which proceeds dominantly through  $S$ -wave annihilation. However, in gaseous hydrogen,  $P$ -wave annihilation is enhanced and, indeed, Ref. [11] reports  $f_1(1420)$  decaying into  $K^*\bar{K}$ . The  $f_1(1420)$ , decaying into  $K\bar{K}\pi$ , is also seen in  $pp$  central production, together with the  $f_1(1285)$ . The

latter decays via  $a_0(980)\pi$ , and the former only via  $K^*\bar{K}$ , while the  $\eta(1440)$  is absent [60,61]. The  $K_S^0 K_S^0 \pi^0$  decay mode of the  $f_1(1420)$  establishes unambiguously  $C=+1$ . On the other hand, there is no evidence for any state decaying into  $\eta\pi\pi$  around 1400 MeV, and hence the  $\eta\pi\pi$  mode of the  $f_1(1420)$  must be suppressed [62].

The COMPASS Collaboration has recently reported an isovector state at 1414 MeV, the  $a_1(1420)$  [63]. This relatively narrow state ( $\simeq 150$  MeV) is produced by diffractive dissociation with 190 GeV pions in  $\pi N \rightarrow 3\pi N$ , decays into  $f_0(980)\pi \rightarrow 3\pi$  (P-wave) and has therefore the quantum numbers  $(I^G)J^{PC} = (1^-)1^{++}$ . The pseudovector nonet already contains the established  $a_1(1260)$  as the  $I = 1$  state. As mentioned above, the  $f_1(1420)$  has been interpreted as a  $K^*\bar{K}$  molecule [51]. The new  $a_1(1420)$  could be its isovector partner. Arguments favoring the  $f_1(1420)$  being a hybrid  $q\bar{q}g$  meson [64] or a four-quark state [65] were also put forward. The  $q\bar{q}$  state would then remain to be identified, with the  $f_1(1510)$  (see below) as a candidate. However, an alternative explanation is suggested in Ref. [66] in which the authors claim a single  $1^{++}$  isovector around 1400 MeV, leading to two peaks in the  $3\pi$  mass spectrum, depending on the production mechanism,  $\rho\pi$  for the  $a_1(1260)$  and  $f_0(980)\pi$  for the  $a_1(1420)$ .

We now turn to the experimental evidence for the  $f_1(1510)$ . The  $f_1(1510)$  was seen in  $K^-p \rightarrow \Lambda K\bar{K}\pi$  at 4 GeV/c [67], and at 11 GeV/c [68]. Evidence is also reported in  $\pi^-p$  at 8 GeV/c, based on the phase motion of the  $1^{++}$   $K^*\bar{K}$  wave [57]. A somewhat broader  $1^{++}$  signal is also observed in  $J/\psi(1S) \rightarrow \gamma\eta\pi^+\pi^-$  [69] as well as a small signal in  $J/\psi(1S) \rightarrow \gamma\eta'\pi^+\pi^-$ , attributed to the  $f_1(1510)$  [70].

The absence of  $f_1(1420)$  in  $K^-p$  [68] argues against the  $f_1(1420)$  being the  $s\bar{s}$  member of the  $1^{++}$  nonet. However, the  $f_1(1420)$  was reported in  $K^-p$  but not in  $\pi^-p$  [71], while two experiments do not observe the  $f_1(1510)$  in  $K^-p$  [71,72]. The latter is also not seen in central collisions [61], nor  $\gamma\gamma$  collisions [73], although, surprisingly for an  $s\bar{s}$  state, a signal is reported in  $4\pi$  decays [74]. These facts led to the conclusion that  $f_1(1510)$  was not well established [75].

Summarizing, there is evidence for two isovector  $1^{++}$  states in the 1400 MeV region, the  $a_1(1260)$  and  $a_1(1420)$ , which cannot be both  $q\bar{q}$  states. These two states could stem from the same pole, or the latter be exotic (tetraquark or hybrid) or a molecular state. The  $f_1(1285)$  and the  $f_1(1420)$  are well known but their nature ( $q\bar{q}$ , tetraquark or molecular) remains to be established. In the  $0^{-+}$  sector there is evidence for two pseudoscalars in the 1400 MeV region, the  $\eta(1405)$  and  $\eta(1475)$ , decaying into  $a_0(980)\pi$  and  $K^*\bar{K}$ , respectively. Alternatively, these two structures could originate from a single pole. Doubts have been expressed on the existence of the  $\eta(1295)$ . The  $f_1(1510)$  remains to be firmly established.

## References

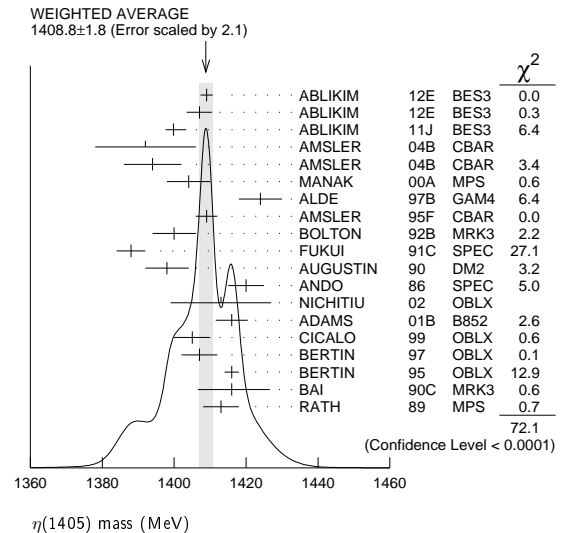
1. P.H. Baillon *et al.*, Nuovo Cimento **50A**, 393 (1967).
2. D.L. Scharre *et al.*, Phys. Lett. **97B**, 329 (1980).

3. C. Edwards *et al.*, Phys. Rev. Lett. **49**, 259 (1982).
4. J.E. Augustin *et al.*, Phys. Rev. **D42**, 10 (1990).
5. J.Z. Bai *et al.*, Phys. Lett. **B594**, 47 (2004).
6. M.G. Rath *et al.*, Phys. Rev. **D40**, 693 (1989).
7. G.S. Adams *et al.*, Phys. Lett. **B516**, 264 (2001).
8. J.Z. Bai *et al.*, Phys. Rev. Lett. **65**, 2507 (1990).
9. J.E. Augustin and G. Cosme, Phys. Rev. **D46**, 1951 (1992).
10. A. Bertin *et al.*, Phys. Lett. **B361**, 187 (1995).
11. A. Bertin *et al.*, Phys. Lett. **B400**, 226 (1997).
12. C. Cicalo *et al.*, Phys. Lett. **B462**, 453 (1999).
13. F. Nichitru *et al.*, Phys. Lett. **B545**, 261 (2002).
14. T. Bolton *et al.*, Phys. Rev. Lett. **69**, 1328 (1992).
15. M. Ablikim *et al.*, Phys. Rev. **D77**, 032005 (2008).
16. M. Ablikim *et al.*, Phys. Rev. Lett. **107**, 182001 (2011).
17. M. Ablikim *et al.*, Phys. Rev. **D87**, 092006 (2013).
18. C. Amsler *et al.*, Phys. Lett. **B358**, 389 (1995).
19. E. Klempt and A. Zaitsev, Phys. Reports **454**, 1 (2007).
20. M. Ablikim *et al.*, Phys. Rev. Lett. **108**, 182001 (2012).
21. J.-J. Wu *et al.*, Phys. Rev. Lett. **108**, 081803 (2012).
22. X.-G. Wu *et al.*, Phys. Rev. **D87**, 014023 (2013).
23. F. Aceti *et al.*, Phys. Rev. **D86**, 114007 (2012).
24. S. Fukui *et al.*, Phys. Lett. **B267**, 293 (1991).
25. D. Alde *et al.*, Phys. Atom. Nucl. **60**, 386 (1997).
26. J.J. Manak *et al.*, Phys. Rev. **D62**, 012003 (2000).
27. A.V. Anisovich *et al.*, Nucl. Phys. **A690**, 567 (2001).
28. A. Abele *et al.*, Phys. Rev. **D57**, 3860 (1998).
29. C. Amsler *et al.*, Eur. Phys. J. **C33**, 23 (2004).
30. B. Aubert *et al.*, Phys. Rev. Lett. **101**, 091801 (2008).
31. E. Klempt, Int. J. Mod. Phys. **A21**, 739 (2006).
32. J. Reinnarth, PhD Thesis, University of Bonn (2003).
33. F. Close *et al.*, Phys. Lett. **B397**, 333 (1997).
34. T. Barnes *et al.*, Phys. Rev. **D55**, 4157 (1997).
35. T. Gutsche *et al.*, Phys. Rev. **D79**, 014036 (2009).
36. M. Acciarri *et al.*, Phys. Lett. **B501**, 1 (2001).
37. F. Close *et al.*, Phys. Rev. **D55**, 5749 (1997).
38. D.M. Li *et al.*, Eur. Phys. J. **C28**, 335 (2003).
39. R. Ahohe *et al.*, Phys. Rev. **D71**, 072001 (2005).
40. P. Achard *et al.*, JHEP **0703**, 018 (2007).
41. G.S. Bali *et al.*, Phys. Lett. **B309**, 378 (1993).
42. C. Morningstar and M. Peardon, Phys. Rev. **D60**, 034509 (1999).
43. L. Faddeev *et al.*, Phys. Rev. **D70**, 114033 (2004).
44. H.-Y. Cheng *et al.*, Phys. Rev. **D79**, 014024 (2009).
45. G. Li *et al.*, J. Phys. **G35**, 055002 (2008).
46. T. Gutsche *et al.*, Phys. Rev. **D80**, 014014 (2009).
47. B. Li, Phys. Rev. **D81**, 114002 (2010).
48. A. Masoni, C. Cicalo, and G.L. Usai, J. Phys. **G32**, R293 (2006).
49. F. Aceti *et al.*, Phys. Lett. **B750**, 609 (2015).
50. S. Stone and L. Zhang, Phys. Rev. Lett. **111**, 062001 (2013).
51. R.S. Longacre, Phys. Rev. **D42**, 874 (1990).
52. R. Aaij *et al.*, Phys. Rev. Lett. **112**, 091802 (2014).
53. G. Gidal *et al.*, Phys. Rev. Lett. **59**, 2012 (1987).

54. C. Dionisi *et al.*, Nucl. Phys. **B169**, 1 (1980).
55. S.U. Chung *et al.*, Phys. Rev. Lett. **55**, 779 (1985).
56. D.F. Reeves *et al.*, Phys. Rev. **D34**, 1960 (1986).
57. A. Birman *et al.*, Phys. Rev. Lett. **61**, 1557 (1988).
58. H.J. Behrend *et al.*, Z. Phys. **C42**, 367 (1989).
59. A. Ando *et al.*, Phys. Rev. Lett. **57**, 1296 (1986).
60. T.A. Armstrong *et al.*, Phys. Lett. **B221**, 216 (1989).
61. D. Barberis *et al.*, Phys. Lett. **B413**, 225 (1997).
62. T.A. Armstrong *et al.*, Z. Phys. **C52**, 389 (1991).
63. C. Adolph *et al.*, Phys. Rev. Lett. **115**, 082001 (2015).
64. S. Ishida *et al.*, Prog. Theor. Phys. **82**, 119 (1989).
65. D.O. Caldwell, *Hadron 89 Conf., Ajaccio, Corsica*, p. 127.
66. J.-L. Basdevant and E.L. Berger, Phys. Rev. Lett. **114**, 192001 (2015).
67. P. Gavillet *et al.*, Z. Phys. **C16**, 119 (1982).
68. D. Aston *et al.*, Phys. Lett. **B201**, 573 (1988).
69. J.Z. Bai *et al.*, Phys. Lett. **B446**, 356 (1999).
70. M. Ablikim *et al.*, Phys. Rev. Lett. **106**, 072002 (2011).
71. S. Bitukov *et al.*, Sov. J. Nucl. Phys. **39**, 738 (1984).
72. E. King *et al.*, Nucl. Phys. (Proc. Supp.) **B21**, 11 (1991).
73. H. Aihara *et al.*, Phys. Rev. **D38**, 1 (1988).
74. D.A. Bauer *et al.*, Phys. Rev. **D48**, 3976 (1993).
75. F.E. Close and A. Kirk, Z. Phys. **C76**, 469 (1997).

 $\eta(1405)$  MASS

VALUE (MeV) DOCUMENT ID  
**1408.8 ± 1.8 OUR AVERAGE** Includes data from the 2 datablocks that follow this one.  
 Error includes scale factor of 2.1. See the ideogram below.

 $\eta\pi\pi$  MODE

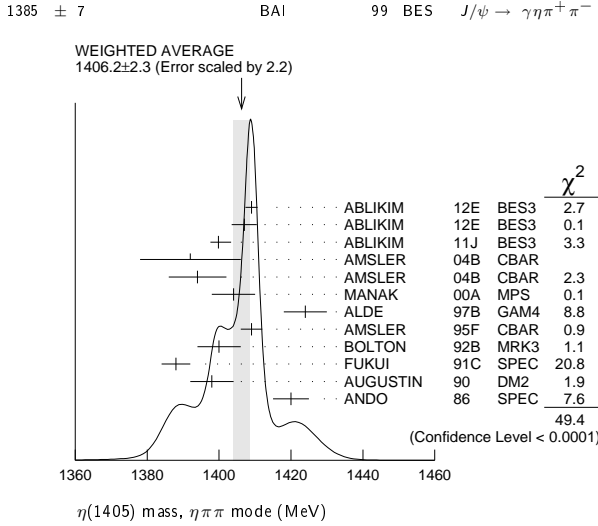
VALUE (MeV) EVTS DOCUMENT ID TECN COMMENT  
 The data in this block is included in the average printed for a previous datablock.

**1406.2 ± 2.3 OUR AVERAGE** Error includes scale factor of 2.2. See the ideogram below.

1409.0 ± 1.7	743	ABLIKIM	12E BES3	$J/\psi \rightarrow \gamma(\pi^+\pi^-\pi^0)\eta$
1407.0 ± 3.5	198	ABLIKIM	12E BES3	$J/\psi \rightarrow \gamma(\pi^0\pi^0\pi^0)\eta$
1399.8 ± 2.2 ± 2.8	1	ABLIKIM	11J BES3	$J/\psi \rightarrow \omega(\eta\pi^+\pi^-)$
1392 ± 14	900 ± 375	AMSLER	04B CBAR	$0\bar{p}p \rightarrow \pi^+\pi^-\pi^+\pi^-\eta$
1394 ± 8	6.6 ± 2.0k	AMSLER	04B CBAR	$0\bar{p}p \rightarrow \pi^+\pi^-\pi^0\pi^0\eta$
1404 ± 6	9082	MANAK	00A MPS	$18\pi^-\pi^0 \rightarrow \eta\pi^+\pi^-\pi^0$
1424 ± 6	2200	ALDE	97B GAM4	$100\pi^-\pi^0 \rightarrow \eta\pi^+\pi^-\pi^0$
1409 ± 3		AMSLER	95F CBAR	$0\bar{p}p \rightarrow \pi^+\pi^-\pi^0\pi^0\eta$
1400 ± 6		BOLTON	92B MRK3	$J/\psi \rightarrow \gamma\eta\pi^+\pi^-$
1388 ± 4		FUKUI	91C SPEC	$8.95\pi^-\pi^0 \rightarrow \eta\pi^+\pi^-\pi^0$
1398 ± 6	261	AUGUSTIN	90 DM2	$J/\psi \rightarrow \gamma\eta\pi^+\pi^-$
1420 ± 5		ANDO	86 SPEC	$8\pi^-\pi^0 \rightarrow \eta\pi^+\pi^-\pi^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

## Meson Particle Listings

 $\eta(1405)$  $K\bar{K}\pi$  MODE ( $a_0(980)\pi$  or direct  $K\bar{K}\pi$ )

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.				
<b>1413.9<math>\pm</math> 1.7 OUR AVERAGE</b>				Error includes scale factor of 1.1.
1413 $\pm$ 14	3651	4 NICHITIU	02 OBLX	
1416 $\pm$ 4 $\pm$ 2	20k	ADAMS	01B B852	18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
1405 $\pm$ 5		5 CICALO	99 OBLX	0 $\bar{p} p \rightarrow K^\pm K_S^0 \pi^\mp \pi^+ \pi^-$
1407 $\pm$ 5		5 BERTIN	97 OBLX	0 $\bar{p} p \rightarrow K^\pm (K^0) \pi^\mp \pi^+ \pi^-$
1416 $\pm$ 2		5 BERTIN	95 OBLX	0 $\bar{p} p \rightarrow K\bar{K} \pi \pi \pi$
1416 $\pm$ 8 $\pm$ 7	700	6 BAI	90C MRK3	$J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
1413 $\pm$ 5		6 RATH	89 MPS	21.4 $\pi^- p \rightarrow n K_S^0 K_S^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1459 $\pm$ 5		7 AUGUSTIN	92 DM2	$J/\psi \rightarrow \gamma K\bar{K} \pi$

 $\pi\pi\gamma$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1390<math>\pm</math>12</b>	235 $\pm$ 91	AMSLER	04B CBAR	0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^+ \pi^- \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1424 $\pm$ 10 $\pm$ 11	547	BAI	04J BES2	$J/\psi \rightarrow \gamma \gamma \pi^+ \pi^-$
1401 $\pm$ 18		8,9 AUGUSTIN	90 DM2	$J/\psi \rightarrow \pi^+ \pi^- \gamma \gamma$
1432 $\pm$ 8		9 COFFMAN	90 MRK3	$J/\psi \rightarrow \pi^+ \pi^- 2\gamma$

 $4\pi$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1420 $\pm$ 20		BUGG	95 MRK3	$J/\psi \rightarrow \gamma \pi^+ \pi^- \pi^+ \pi^-$
1489 $\pm$ 12	3270	10 BISELLO	89B DM2	$J/\psi \rightarrow 4\pi \gamma$

 $K\bar{K}\pi$  MODE (unresolved)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1452.7 $\pm$ 3.3	191	11,12 ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K K \pi$
1437.6 $\pm$ 3.2	249 $\pm$ 35	11,12 ABLIKIM	08E BES2	$J/\psi \rightarrow \omega K_S^0 K^+ \pi^- + \text{c.c.}$
1445.9 $\pm$ 5.7	62 $\pm$ 18	11,12 ABLIKIM	08E BES2	$J/\psi \rightarrow \omega K^+ K^- \pi^0$
1442 $\pm$ 10	410	11 BAI	98C BES	$J/\psi \rightarrow \gamma K^+ K^- \pi^0$
1445 $\pm$ 8	693	11 AUGUSTIN	90 DM2	$J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
1433 $\pm$ 8	296	11 AUGUSTIN	90 DM2	$J/\psi \rightarrow \gamma K^+ K^- \pi^0$
1413 $\pm$ 8	500	11 DUCH	89 ASTE	$\bar{p} p \rightarrow \pi^+ \pi^- K^\pm \pi^\mp K^0$
1453 $\pm$ 7	170	11 RATH	89 MPS	21.4 $\pi^- p \rightarrow K_S^0 K_S^0 \pi^0 n$
1419 $\pm$ 1	8800	11 BIRMAN	88 MPS	8 $\pi^- p \rightarrow K^+ \bar{K}^0 \pi^- n$
1424 $\pm$ 3	620	11 REEVES	86 SPEC	6.6 $p\bar{p} \rightarrow K\bar{K} \pi X$
1421 $\pm$ 2		11 CHUNG	85 SPEC	8 $\pi^- p \rightarrow K\bar{K} \pi n$
1440 $\pm$ 20	174	11 EDWARDS	82E CBAL	$J/\psi \rightarrow \gamma K^+ K^- \pi^0$
1440 $\pm$ 10		11 SCHARRE	80 MRK2	$J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
1425 $\pm$ 7	800	11,13 BAILLON	67 HBC	0 $\bar{p} p \rightarrow K\bar{K} \pi \pi \pi$

- The selected process is  $J/\psi \rightarrow \omega a_0(980) \pi$ .
- From fit to the  $a_0(980) \pi$   $0^-+$  partial wave.
- Best fit with a single Breit Wigner.
- Decaying dominantly directly to  $K^+ K^- \pi^0$ .
- Decaying into  $(K\bar{K})_S \pi$ ,  $(K\pi)_S \bar{K}$ , and  $a_0(980) \pi$ .
- From fit to the  $a_0(980) \pi$   $0^-+$  partial wave. Cannot rule out a  $a_0(980) \pi$   $1^-+$  partial wave.
- Excluded from averaging because averaging would be meaningless.
- Best fit with a single Breit Wigner.
- This peak in the  $\gamma \rho$  channel may not be related to the  $\eta(1405)$ .

<sup>10</sup> Estimated by us from various fits.

<sup>11</sup> These experiments identify only one pseudoscalar in the 1400–1500 range. Data could also refer to  $\eta(1475)$ .

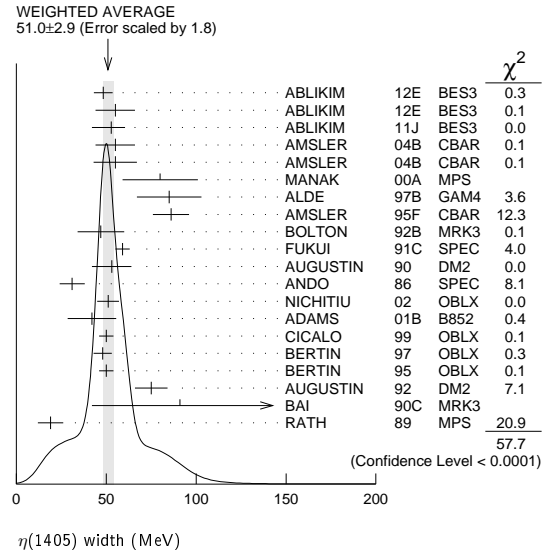
<sup>12</sup> Systematic uncertainty not evaluated.

<sup>13</sup> From best fit of  $0^-+$  partial wave, 50%  $K^*(892) K$ , 50%  $a_0(980) \pi$ .

 $\eta(1405)$  WIDTH

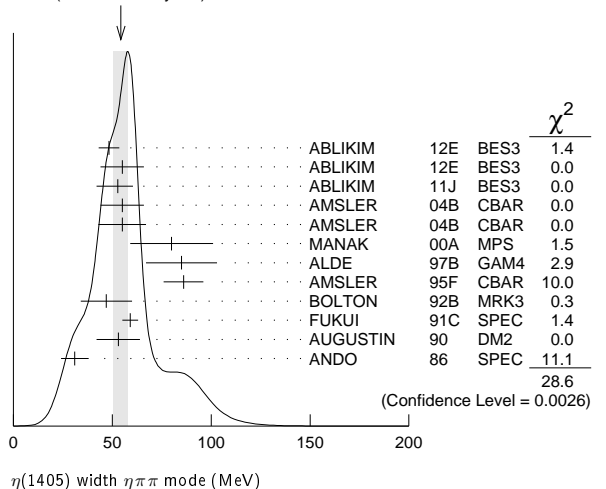
VALUE (MeV)      DOCUMENT ID

**51.0 $\pm$ 2.9 OUR AVERAGE** Includes data from the 2 datablocks that follow this one. Error includes scale factor of 1.8. See the ideogram below.

 $\eta\pi\pi$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.				
<b>54 <math>\pm</math> 4 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
48.3 $\pm$ 5.2	743	ABLIKIM	12E BES3	$J/\psi \rightarrow \gamma (\pi^+ \pi^- \pi^0)$
55.0 $\pm$ 11.0	198	ABLIKIM	12E BES3	$J/\psi \rightarrow \gamma (\pi^0 \pi^0 \pi^0)$
52.8 $\pm$ 7.6 $\pm$ 0.1		14 ABLIKIM	11J BES3	$J/\psi \rightarrow \omega (\eta \pi^+ \pi^-)$
55 $\pm$ 11	900 $\pm$ 375	AMSLER	04B CBAR	0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^+ \pi^- \eta$
55 $\pm$ 12	6.6 $\pm$ 2.0k	AMSLER	04B CBAR	0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^0 \pi^0 \gamma$
80 $\pm$ 21	9082	MANAK	00A MPS	18 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
85 $\pm$ 18	2200	ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta \pi^0 \pi^0 n$
86 $\pm$ 10		AMSLER	95F CBAR	0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^0 \pi^0 \eta$
47 $\pm$ 13		15 BOLTON	92B MRK3	$J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$
59 $\pm$ 4		FUKUI	91C SPEC	8.95 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
53 $\pm$ 11		16 AUGUSTIN	90 DM2	$J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$
31 $\pm$ 7		ANDO	86 SPEC	8 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$

WEIGHTED AVERAGE  
54 $\pm$ 4 (Error scaled by 1.6)



See key on page 601

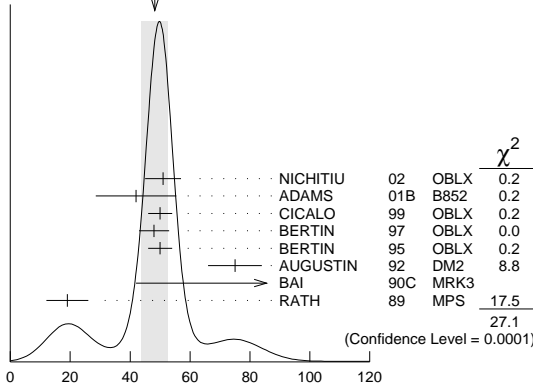
## Meson Particle Listings

 $\eta(1405)$  $K\bar{K}\pi$  MODE ( $a_0(980)\pi$  or direct  $K\bar{K}\pi$ )

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

<b>48±4 OUR AVERAGE</b>	Error includes scale factor of 2.1. See the ideogram below.			
51±6	3651	17 NICHITIU	02 OBLX	
42±10±9	20k	ADAMS	01B B852	18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
50±4		CICALO	99 OBLX	0 $\bar{p} p \rightarrow K^\pm K_S^0 \pi^\mp \pi^\pm \pi^-$
48±5		18 BERTIN	97 OBLX	0.0 $\bar{p} p \rightarrow K^\pm (K^0) \pi^\mp \pi^\pm \pi^-$
50±4		18 BERTIN	95 OBLX	0 $\bar{p} p \rightarrow K\bar{K}\pi\pi\pi$
75±9		AUGUSTIN	92 DM2	$J/\psi \rightarrow \gamma K\bar{K}\pi$
91+67+15 -31-38		19 BAI	90c MRK3	$J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
19±7		19 RATH	89 MPS	21.4 $\pi^- p \rightarrow n K_S^0 K_S^0 \pi^0$

WEIGHTED AVERAGE  
48±4 (Error scaled by 2.1) $\eta(1405)$  width  $K\bar{K}\pi$  mode ( $a_0(980)\pi$  dominant) $\pi\pi\gamma$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

101.0±8.8±8.8	547	BAI	04J BES2	$J/\psi \rightarrow \gamma\gamma\pi^+\pi^-$
174±44		AUGUSTIN	90 DM2	$J/\psi \rightarrow \pi^+\pi^-\gamma\gamma$
90±26	20	COFFMAN	90 MRK3	$J/\psi \rightarrow \pi^+\pi^-2\gamma$

 $4\pi$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

160±30		BUGG	95 MRK3	$J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
144±13	3270	21 BISELLO	89B DM2	$J/\psi \rightarrow 4\pi\gamma$

 $K\bar{K}\pi$  MODE (unresolved)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

45.9±8.2	191	22,23 ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K K \pi$
48.9±9.0	249±35	22,23 ABLIKIM	08E BES2	$J/\psi \rightarrow \omega K_S^0 K^+\pi^- + c.c.$
34.2±18.5	62±18	22,23 ABLIKIM	08E BES2	$J/\psi \rightarrow \omega K^+ K^- \pi^0$
93±14	296	22 AUGUSTIN	90 DM2	$J/\psi \rightarrow \gamma K^+ K^- \pi^0$
105±10	693	22 AUGUSTIN	90 DM2	$J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
62±16	500	22 DUCH	89 ASTE	$\bar{p} p \rightarrow K\bar{K}\pi\pi\pi$
100±11	170	22 RATH	89 MPS	21.4 $\pi^- p \rightarrow K_S^0 K_S^0 \pi^0 n$
66±2	8800	22 BIRMAN	88 MPS	8 $\pi^- p \rightarrow K^+ \bar{K}^0 \pi^- n$
60±10	620	22 REEVES	86 SPEC	6.6 $p\bar{p} \rightarrow K K \pi X$
60±10		22 CHUNG	85 SPEC	8 $\pi^- p \rightarrow K\bar{K}\pi n$
55+20 -30	174	22 EDWARDS	82E CBAL	$J/\psi \rightarrow \gamma K^+ K^- \pi^0$
50+30 -20		22 SCHARRE	80 MRK2	$J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
80±10	800	22,24 BAILLON	67 HBC	0.0 $\bar{p} p \rightarrow K\bar{K}\pi\pi\pi$

14 The selected process is  $J/\psi \rightarrow \omega a_0(980)\pi$ .15 From fit to the  $a_0(980)\pi 0^-+$  partial wave.16 From  $\eta\pi^+\pi^-$  mass distribution - mainly  $a_0(980)\pi$  - no spin-parity determination available.17 Decaying dominantly directly to  $K^+ K^- \pi^0$ .18 Decaying into  $(K\bar{K})_S \pi$ ,  $(K\pi)_S \bar{K}$ , and  $a_0(980)\pi$ .19 From fit to the  $a_0(980)\pi 0^-+$  partial wave, but  $a_0(980)\pi 1^++$  cannot be excluded.20 This peak in the  $\gamma\rho$  channel may not be related to the  $\eta(1405)$ .

21 Estimated by us from various fits.

22 These experiments identify only one pseudoscalar in the 1400–1500 range. Data could also refer to  $\eta(1475)$ .

23 Systematic uncertainty not evaluated.

24 From best fit to  $0^-+$  partial wave, 50%  $K^*(892)K$ , 50%  $a_0(980)\pi$ . $\eta(1405)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
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$\Gamma_1$ $K\bar{K}\pi$	seen	
$\Gamma_2$ $\eta\pi\pi$	seen	
$\Gamma_3$ $a_0(980)\pi$	seen	
$\Gamma_4$ $\eta(\pi\pi)s$ -wave	seen	
$\Gamma_5$ $f_0(980)\eta$	seen	
$\Gamma_6$ $4\pi$	seen	
$\Gamma_7$ $\rho\rho$	<58 %	99.85%
$\Gamma_8$ $\gamma\gamma$		
$\Gamma_9$ $\rho^0\gamma$	seen	
$\Gamma_{10}$ $\phi\gamma$		
$\Gamma_{11}$ $K^*(892)K$	seen	

 $\eta(1405)$   $\Gamma(i)/\Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

$\Gamma(K\bar{K}\pi) \times \Gamma(\gamma\gamma)/\Gamma(\text{total})$	VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.035	90	25,26	AHOHE	05	CLE2 10.6 $e^+e^- \rightarrow e^+e^- K_S^0 K^\pm \pi^\mp$
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 $\Gamma(\eta\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
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<0.095 95 ACCIARRI 01G L3 183–202  $e^+e^- \rightarrow e^+e^-\eta\pi^+\pi^-$

 $\Gamma(\rho^0\gamma) \times \Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.5	95	ALTHOFF	84E TASS	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-\gamma$
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25 Using  $\eta(1405)$  mass and width 1410 MeV and 51 MeV, respectively.  
26 Assuming three-body phase-space decay to  $K_S^0 K^\pm \pi^\mp$ .

 $\eta(1405)$  BRANCHING RATIOS

$\Gamma(\eta\pi\pi)/\Gamma(K\bar{K}\pi)$	VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.09±0.48	27	AMSLER	04B CBAR	0 $\bar{p} p \rightarrow \pi^+\pi^-\pi^+\pi^- \eta$
<0.5	90	EDWARDS	83B CBAL	$J/\psi \rightarrow \eta\pi\pi\gamma$
<1.1	90	SCHARRE	80 MRK2	$J/\psi \rightarrow \eta\pi\pi\gamma$
<1.5	95	FOSTER	68B HBC	0.0 $\bar{p} p$

$\Gamma(\rho^0\gamma)/\Gamma(\eta\pi\pi)$	VALUE	DOCUMENT ID	TECN	COMMENT
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0.111±0.064 AMSLER 04B CBAR 0  $\bar{p} p$

$\Gamma(a_0(980)\pi)/\Gamma(K\bar{K}\pi)$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

~0.15	28	BERTIN	95 OBLX	0 $\bar{p} p \rightarrow K\bar{K}\pi\pi\pi$
~0.8	500	28 DUCH	89 ASTE	$\bar{p} p \rightarrow \pi^+\pi^- K^\pm \pi^\mp K^0$
~0.75	28	REEVES	86 SPEC	6.6 $p\bar{p} \rightarrow K K \pi X$

$\Gamma(a_0(980)\pi)/\Gamma(\eta\pi\pi)$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.29±0.10	2200	ABELE	98E CBAR	0 $p\bar{p} \rightarrow \eta\pi^0\pi^0\pi^0$
0.19±0.04	29	ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta\pi^0\pi^0 n$
0.56±0.04±0.03	29	AMSLER	95F CBAR	0 $\bar{p} p \rightarrow \pi^+\pi^-\pi^0\pi^0\eta$

$\Gamma(a_0(980)\pi)/\Gamma(\eta(\pi\pi)_{S\text{-wave}})$					$\Gamma_3/\Gamma_4$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.91 \pm 0.12$		ANISOVICH	01	SPEC $0.0 \bar{p} p \rightarrow \eta \pi^+ \pi^- \pi^+ \pi^-$	
$0.15 \pm 0.04$	9082	30 MANAK	00A	MPS $18 \pi^- p \rightarrow \eta \pi^+ \pi^- n$	
$0.70 \pm 0.12 \pm 0.20$		31 BAI	99	BES $J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$	

$\Gamma(\rho^0\gamma)/\Gamma(K\bar{K}\pi)$	VALUE	DOCUMENT ID	TECN	COMMENT
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0.0152±0.0038 32 COFFMAN 90 MRK3  $J/\psi \rightarrow \gamma\gamma\pi^+\pi^-$

$\Gamma(\eta(\pi\pi)s\text{-wave})/\Gamma(\eta\pi\pi)$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.81±0.04	2200	ALDE	97B GAM4	100 $\pi^- p \rightarrow \eta\pi^0\pi^0 n$
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# Meson Particle Listings

## $\eta(1405)$ , $f_1(1420)$

$\Gamma(\bar{\eta}_0(980)\eta)/\Gamma(\eta\pi\pi)$				$\Gamma_5/\Gamma_2$	
VALUE		DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.32 \pm 0.07$		33 ANISOVICH	00	SPEC	$0.9\text{--}1.2 \bar{p}p \rightarrow \eta 3\pi^0$
$\Gamma(\rho\rho)/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.58	99.85	27,34 AMSLER	04B	CBAR	0 $\bar{p}p$
$\Gamma(K^*(892)K)/\Gamma(a_0(980)\pi)$				$\Gamma_{11}/\Gamma_3$	
VALUE		DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.084 \pm 0.024$		30 ADAMS	01B	B852	18 GeV $\pi^-p \rightarrow K^+K^-\pi^0 n$
$\Gamma(\phi\gamma)/\Gamma(\rho^0\gamma)$				$\Gamma_{10}/\Gamma_9$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.77	95	35 BAI	04J	BES2	$J/\psi \rightarrow \gamma\gamma K^+K^-$
27 Using the data of BAILLON 67 on $\bar{p}p \rightarrow K\bar{K}\pi$ .					
28 Assuming that the $a_0(980)$ decays only into $K\bar{K}$ .					
29 Assuming that the $a_0(980)$ decays only into $\eta\pi$ .					
30 Statistical error only.					
31 Assuming that the $a_0(980)$ decays only into $\eta\pi$ .					
32 Using $B(J/\psi \rightarrow \gamma\eta(1405) \rightarrow \gamma K\bar{K}\pi) = 4.2 \times 10^{-3}$ and $B(J/\psi \rightarrow \gamma\eta(1405) \rightarrow \gamma\gamma\rho^0) = 6.4 \times 10^{-5}$ and assuming that the $\gamma\rho^0$ signal does not come from the $f_1(1420)$ .					
33 Using preliminary Crystal Barrel data.					
34 Assuming that the $\eta(1405)$ decays are saturated by the $\pi\pi\eta$ , $K\bar{K}\pi$ and $\rho\rho$ modes.					
35 Calculated by us from $B(J/\psi \rightarrow \eta(1405)\gamma \rightarrow \phi\gamma\gamma) < 0.82 \times 10^{-4}$ and $B(J/\psi \rightarrow \eta(1405)\gamma \rightarrow \rho^0\gamma\gamma) = (1.07 \pm 0.17 \pm 0.11) \times 10^{-4}$ .					

### $\eta(1405)$ REFERENCES

ABLIKIM	13M	PR D87 092006	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12E	PRL 108 182001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11J	PRL 107 182001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	08E	PR D77 032005	M. Ablikim <i>et al.</i>	(BES Collab.)
AHOHE	05	PR D71 072001	R. Ahohe <i>et al.</i>	(CLEO Collab.)
AMSLER	04B	EPJ C33 23	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
BAI	04J	PL B594 47	J.Z. Bai <i>et al.</i>	(BES Collab.)
NICHITIU	02	PL B545 261	F. Nichitiu <i>et al.</i>	(OBELIX Collab.)
ACCIARRI	01G	PL B501 1	M. Acciarri <i>et al.</i>	(L3 Collab.)
ADAMS	01B	PL B516 264	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)
ANISOVICH	01	NP A690 567	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00	PL B472 168	A.V. Anisovich <i>et al.</i>	
MANAK	00A	PR D62 012003	J.J. Manak <i>et al.</i>	(BNL E852 Collab.)
BAI	99	PL B446 356	J.Z. Bai <i>et al.</i>	(BES Collab.)
CICALO	99	PL B462 453	C. Cicalo <i>et al.</i>	(OBELIX Collab.)
ABELE	98E	NP B514 45	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BAI	98C	PL B440 217	J.Z. Bai <i>et al.</i>	(BES Collab.)
ALDE	97B	PAN 60 386	D. Alde <i>et al.</i>	(GAMS Collab.)
Translated from YAF 60 458.				
BERTIN	97	PL B400 226	A. Bertin <i>et al.</i>	(OBELIX Collab.)
AMSLER	95F	PL B358 389	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
BERTIN	95	PL B361 187	A. Bertin <i>et al.</i>	(OBELIX Collab.)
BUGG	95	PL B353 378	D.V. Bugg <i>et al.</i>	(LOQM, PNPI, WASH)
AUGUSTIN	92	PR D46 1951	J.E. Augustin, G. Cosme	(DM2 Collab.)
BOLTON	92B	PRL 69 1328	T. Bolton <i>et al.</i>	(Mark III Collab.)
FUKUI	91C	PL B267 293	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
AUGUSTIN	90	PR D42 10	J.E. Augustin <i>et al.</i>	(DM2 Collab.)
BAI	90C	PRL 65 2507	Z. Bai <i>et al.</i>	(Mark III Collab.)
COFFMAN	90	PR D41 1410	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
BISELLO	89B	PR D39 701	G. Busetto <i>et al.</i>	(DM2 Collab.)
DUCH	89	ZPHY C45 223	K.D. Duch <i>et al.</i>	(ASTERIX Collab.)
RATH	89	PR D40 693	M.G. Rath <i>et al.</i>	(NDAM, BRAN, BNL, CUNY+)
BIRMAN	88	PRL 61 1557	A. Birman <i>et al.</i>	(BNL, FSU, IND, MASD)
ANDO	86	PRL 57 1296	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
REEVES	86	PR D34 1960	D.F. Reeves <i>et al.</i>	(FLOR, BNL, IND+)
CHUNG	85	PRL 55 779	S.U. Chung <i>et al.</i>	(BNL, FLOR, IND+)
ALTHOFF	84E	PL 147B 487	M. Althoff <i>et al.</i>	(TASSO Collab.)
EDWARDS	83B	PRL 51 859	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
EDWARDS	82E	PRL 49 259	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
Also				
SCHARRE	80	PL 97B 329	D.L. Scharre <i>et al.</i>	(CIT, HARV, PRIN+)
FOSTER	68B	NP B8 174	M. Foster <i>et al.</i>	(SLAC, LBL)
BAILLON	67	NC 50A 393	P.H. Baillon <i>et al.</i>	(CERN, CDEF)

## $f_1(1420)$

$$I^G(J^{PC}) = 0^+(1^{++})$$

See the minireview under  $\eta(1405)$ .

### $f_1(1420)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
Error includes scale factor of 1.1.				
1434 ± 5 ± 5	133	1 ACHARD	07	L3
183–209 $e^+e^- \rightarrow e^+e^-K_S^0K^\pm\pi^\mp$				
1426 ± 6	711	ABDALLAH	03H	DLPH
91.2 $e^+e^- \rightarrow K_S^0K^\pm\pi^\mp + X$				
1420 ± 14	3651	NICHITIU	02	OBLX
1428 ± 4 ± 2	20k	ADAMS	01B	B852
18 GeV $\pi^-p \rightarrow K^+K^-\pi^0 n$				
1426 ± 1		BARBERIS	97c	OMEG
450 $pp \rightarrow ppK_S^0K^\pm\pi^\mp$				
1425 ± 8		BERTIN	97	OBLX
0.0 $\bar{p}p \rightarrow K^\pm(K^0)\pi^\mp\pi^+\pi^-$				
1435 ± 9		PROKOSHKIN	97B	GAM4
100 $\pi^-p \rightarrow \eta\pi^0\pi^0 n$				

1430 ± 4		2 ARMSTRONG	92E	OMEG	85,300 $\pi^+p, pp \rightarrow \pi^+p, pp(K\bar{K}\pi)$
1462 ± 20		3 AUGUSTIN	92	DM2	$J/\psi \rightarrow \gamma K\bar{K}\pi$
1443 ± 7 ± 3 ± 2	1100	BAI	90C	MRK3	$J/\psi \rightarrow \gamma K_S^0K^\pm\pi^\mp$
1425 ± 10	17	BEHREND	89	CELL	$\gamma\gamma \rightarrow K_S^0K^\pm\pi^\mp$
1442 ± 5 ± 10 ± 17	111	BECKER	87	MRK3	$e^+e^- \rightarrow \omega K\bar{K}\pi$
1423 ± 4		GIDAL	87B	MRK2	$e^+e^- \rightarrow e^+e^-K\bar{K}\pi$
1417 ± 13	13	AIHARA	86c	TPC	$e^+e^- \rightarrow e^+e^-K\bar{K}\pi$
1422 ± 3		CHAUVAT	84	SPEC	ISR 31.5 $pp$
1440 ± 10		4 BROMBERG	80	SPEC	100 $\pi^-p \rightarrow K\bar{K}\pi X$
1426 ± 6	221	DIONISI	80	HBC	4 $\pi^-p \rightarrow K\bar{K}\pi n$
1420 ± 20		DAHL	67	HBC	1.6–4.2 $\pi^-p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1430.8 ± 0.9		5 SOSA	99	SPEC	$pp \rightarrow p_{\text{slow}}(K_S^0K^+\pi^-) p_{\text{fast}}$
1433.4 ± 0.8		5 SOSA	99	SPEC	$pp \rightarrow p_{\text{slow}}(K_S^0K^-\pi^+) p_{\text{fast}}$
1429 ± 3	389	ARMSTRONG	89	OMEG	300 $pp \rightarrow K\bar{K}\pi pp$
1425 ± 2	1520	ARMSTRONG	84	OMEG	85 $\pi^+p, pp \rightarrow (\pi^+,p)(K\bar{K}\pi)p$
~ 1420		BITYUKOV	84	SPEC	32 $K^-p \rightarrow K^+K^-\pi^0\gamma$

- 1 From a fit with a width fixed at 55 MeV.  
2 This result supersedes ARMSTRONG 84, ARMSTRONG 89.  
3 From fit to the  $K^*(892)K1^{++}$  partial wave.  
4 Mass error increased to account for  $a_0(980)$  mass cut uncertainties.  
5 No systematic error given.

### $f_1(1420)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
54.9 ± 2.6 OUR AVERAGE				
51 ± 14	711	ABDALLAH	03H	DLPH 91.2 $e^+e^- \rightarrow K_S^0 K^\pm \pi^\mp + X$
61 ± 8	3651	NICHITIU	02	OBLX
38 ± 9 ± 6	20k	ADAMS	01B	B852 18 GeV $\pi^- p \rightarrow K^+ K^- \pi^0 n$
58 ± 4		BARBERIS	97c	OMEG 450 $pp \rightarrow pp K_S^0 K^\pm \pi^\mp$
45 ± 10		BERTIN	97	OBLX 0.0 $\bar{p}p \rightarrow K^\pm (K^0) \pi^\mp \pi^+ \pi^-$
90 ± 25		PROKOSHKIN	97B	GAM4 100 $\pi^- p \rightarrow \eta \pi^0 \pi^0 n$
58 ± 10		6 ARMSTRONG	92E	OMEG 85,300 $\pi^+ p, pp \rightarrow \pi^+ p, pp (K \bar{K} \pi)$
129 ± 41		7 AUGUSTIN	92	DM2 $J/\psi \rightarrow \gamma K \bar{K} \pi$
68 $\begin{smallmatrix} +29 \\ -18 \\ -9 \end{smallmatrix}$	1100	BAI	90C	MRK3 $J/\psi \rightarrow \gamma K_S^0 K^\pm \pi^\mp$
42 ± 22	17	BEHREND	89	CELL $\gamma \gamma \rightarrow K_S^0 K^\pm \pi^\mp$
40 $\begin{smallmatrix} +17 \\ -13 \end{smallmatrix}$ ± 5	111	BECKER	87	MRK3 $e^+ e^- \rightarrow \omega K \bar{K} \pi$
35 $\begin{smallmatrix} +47 \\ -20 \end{smallmatrix}$	13	AIHARA	86c	TPC $e^+ e^- \rightarrow e^+ e^- K \bar{K} \pi$
47 ± 10		CHAUVAT	84	SPEC ISR 31.5 $pp$
62 ± 14		BROMBERG	80	SPEC 100 $\pi^- p \rightarrow K \bar{K} \pi X$
40 ± 15	221	DIONISI	80	HBC 4 $\pi^- p \rightarrow K \bar{K} \pi n$
60 ± 20		DAHL	67	HBC 1.6-4.2 $\pi^- p$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
68.7 ± 2.9		8 SOSA	99	SPEC $pp \rightarrow p_{\text{slow}} (K_S^0 K^+ \pi^-) p_{\text{fast}}$
58.8 ± 3.3		8 SOSA	99	SPEC $pp \rightarrow p_{\text{slow}} (K_S^0 K^- \pi^+) p_{\text{fast}}$
58 ± 8	389	ARMSTRONG	89	OMEG 300 $pp \rightarrow K \bar{K} \pi pp$
62 ± 5	1520	ARMSTRONG	84	OMEG 85 $\pi^+ p, pp \rightarrow (\pi^+, p) (K \bar{K} \pi) p$
~ 50		BITYUKOV	84	SPEC 32 $K^- p \rightarrow K^+ K^- \pi^0 \gamma$

- 6 This result supersedes ARMSTRONG 84, ARMSTRONG 89.  
7 From fit to the  $K^*(892)K1^{++}$  partial wave.  
8 No systematic error given.

### $f_1(1420)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\bar{K}\pi$	dominant
$\Gamma_2$ $K\bar{K}^*(892) + \text{c.c.}$	dominant
$\Gamma_3$ $\eta\pi\pi$	possibly seen
$\Gamma_4$ $a_0(980)\pi$	
$\Gamma_5$ $\pi\pi\rho$	
$\Gamma_6$ $4\pi$	
$\Gamma_7$ $\rho^0\gamma$	
$\Gamma_8$ $\phi\gamma$	seen



See key on page 601

# Meson Particle Listings

## $f_1(1420), \omega(1420)$

$f_1(1420) \Gamma(i) \Gamma(\gamma\gamma) / \Gamma(\text{total})$					
$\Gamma(K\bar{K}\pi) \times \Gamma(\gamma\gamma^*) / \Gamma(\text{total})$					
VALUE (keV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.9 \pm 0.4</math> OUR AVERAGE</b>					
$3.2 \pm 0.6 \pm 0.7$	133	<sup>9,10</sup>	ACHARD	07 L3	$183\text{--}209 e^+ e^- \rightarrow e^+ e^- K_S^0 K^\pm \pi^\mp$
$3.0 \pm 0.9 \pm 0.7$		<sup>11,12</sup>	BEHREND	89 CELL	$e^+ e^- \rightarrow e^+ e^- K_S^0 K^\pm \pi^\mp$
$2.3^{+1.0}_{-0.9} \pm 0.8$			HILL	89 JADE	$e^+ e^- \rightarrow e^+ e^- K^\pm K_S^0 \pi^\mp$
$1.3 \pm 0.5 \pm 0.3$			AIHARA	88b TPC	$e^+ e^- \rightarrow e^+ e^- K^\pm K_S^0 \pi^\mp$
$1.6 \pm 0.7 \pm 0.3$		<sup>11,13</sup>	GIDAL	87b MRK2	$e^+ e^- \rightarrow e^+ e^- K^\pm K_S^0 \pi^\mp$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<8.0	95		JENNI	83 MRK2	$e^+ e^- \rightarrow e^+ e^- K^\pm K^\mp \pi^\mp$

$f_1(1420) \text{ BRANCHING RATIOS}$					
$\Gamma(K\bar{K}^*(892) + \text{c.c.}) / \Gamma(K\bar{K}\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_1$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.76 \pm 0.06$		BROMBERG	80 SPEC	$100 \pi^- p \rightarrow K\bar{K}\pi X$	
$0.86 \pm 0.12$		DIONISI	80 HBC	$4 \pi^- p \rightarrow K\bar{K}\pi n$	

$\Gamma(\pi\pi\rho) / \Gamma(K\bar{K}\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma_1$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.3	95	CORDEN	78 OMEG	$12\text{--}15 \pi^- p$	
<2.0		DAHL	67 HBC	$1.6\text{--}4.2 \pi^- p$	

$\Gamma(\eta\pi\pi) / \Gamma(K\bar{K}\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma_1$
<b>&lt;0.1</b>	95	ARMSTRONG	91b OMEG	$300 pp \rightarrow pp\pi^+\pi^-\pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.35 \pm 0.75$		KOPKE	89 MRK3	$J/\psi \rightarrow \omega\eta\pi\pi(K\bar{K}\pi)$	
<0.6	90	GIDAL	87 MRK2	$e^+ e^- \rightarrow e^+ e^- \eta\pi^+\pi^-$	
<0.5	95	CORDEN	78 OMEG	$12\text{--}15 \pi^- p$	
$1.5 \pm 0.8$		DEFOIX	72 HBC	$0.7 \bar{p} p$	

$\Gamma(a_0(980)\pi) / \Gamma(\eta\pi\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma_3$
<b>&gt;0.1</b>	90	PROKOSHKIN	97b GAM4	$100 \pi^- p \rightarrow \eta\pi^0\pi^0 n$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
not seen in either mode		ANDO	86 SPEC	$8 \pi^- p$	
not seen in either mode		CORDEN	78 OMEG	$12\text{--}15 \pi^- p$	
$0.4 \pm 0.2$		DEFOIX	72 HBC	$0.7 \bar{p} p \rightarrow 7\pi$	

$\Gamma(4\pi) / \Gamma(K\bar{K}^*(892) + \text{c.c.})$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma_2$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.90	95	DIONISI	80 HBC	$4 \pi^- p$	

$\Gamma(K\bar{K}\pi) / [\Gamma(K\bar{K}^*(892) + \text{c.c.}) + \Gamma(a_0(980)\pi)]$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/(\Gamma_2+\Gamma_4)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.65 \pm 0.27$		<sup>14</sup> DIONISI	80 HBC	$4 \pi^- p$	
<sup>14</sup> Calculated using $\Gamma(K\bar{K})/\Gamma(\eta\pi) = 0.24 \pm 0.07$ for $a_0(980)$ fractions.					

$\Gamma(a_0(980)\pi) / \Gamma(K\bar{K}^*(892) + \text{c.c.})$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma_2$
<b><math>0.04 \pm 0.01 \pm 0.01</math></b>		BARBERIS	98c OMEG	$450 pp \rightarrow \rho_f f_1(1420) p_S$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.04	68	ARMSTRONG	84 OMEG	$85 \pi^+ p$	

$\Gamma(4\pi) / \Gamma(K\bar{K}\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma_1$
<b>&lt;0.62</b>	95	ARMSTRONG	89g OMEG	$85 \pi p \rightarrow 4\pi X$	

$\Gamma(\rho^0\gamma) / \Gamma_{\text{total}}$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
<0.08	95	<sup>15</sup> ARMSTRONG	92c SPEC	$300 pp \rightarrow pp\pi^+\pi^-\gamma$	
<sup>15</sup> Using the data on the $\bar{K} K \pi$ mode from ARMSTRONG 89.					

$\Gamma(\rho^0\gamma) / \Gamma(K\bar{K}\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma_1$
<b>&lt;0.02</b>	95	BARBERIS	98c OMEG	$450 pp \rightarrow \rho_f f_1(1420) p_S$	

$\Gamma(\phi\gamma) / \Gamma(K\bar{K}\pi)$					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma_1$
<b><math>0.003 \pm 0.001 \pm 0.001</math></b>		BARBERIS	98c OMEG	$450 pp \rightarrow \rho_f f_1(1420) p_S$	

## $f_1(1420)$ REFERENCES

ACHARD	07	JHEP 0703 018	P. Achard <i>et al.</i>	(L3 Collab.)
ABDALLAH	03H	PL B569 129	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
NICHITU	02	PL B545 261	F. Nichitiu <i>et al.</i>	(OBELIX Collab.)
ADAMS	01B	PL B516 264	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)
SOSA	99	PRL 83 913	M. Sosa <i>et al.</i>	
BARBERIS	98C	PL B440 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	97C	PL B413 225	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BERTIN	97	PL B400 226	A. Bertin <i>et al.</i>	(OBELIX Collab.)
PROKOSHKIN	97B	SPD 42 298	Yu.D. Prokoshkin, S.A. Sadovsky	
Translated from DANS 354 751.				
ARMSTRONG	92C	ZPHY C54 371	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
ARMSTRONG	92E	ZPHY C56 29	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+) JPC
AUGUSTIN	92	PR D46 1951	J.E. Augustin, G. Cosme	(DM2 Collab.)
ARMSTRONG	91B	ZPHY C52 389	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+)
BAI	90C	PRL 65 2507	Z. Bai <i>et al.</i>	(Mark III Collab.)
ARMSTRONG	89	PL B221 216	T.A. Armstrong <i>et al.</i>	(CERN, CDEF, BIRM+) JPC
ARMSTRONG	89G	ZPHY C43 55	T.A. Armstrong <i>et al.</i>	(CERN, BIRM, BARI+)
BEHREND	89	ZPHY C42 367	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
HILL	89	ZPHY C42 355	P. Hill <i>et al.</i>	(JADE Collab.) JP
KOPKE	89	PRPL 174 67	L. Kopke <i>et al.</i>	(CERN)
AIHARA	88B	PL B209 107	H. Aihara <i>et al.</i>	(TPC-2 $\gamma$ Collab.)
BECKER	87	PRL 59 186	J.J. Becker <i>et al.</i>	(Mark III Collab.) JP
GIDAL	87	PRL 59 2012	G. Gidal <i>et al.</i>	(LBL, SLAC, HARV)
GIDAL	87B	PRL 59 2016	G. Gidal <i>et al.</i>	(LBL, SLAC, HARV)
AIHARA	86C	PRL 57 2500	H. Aihara <i>et al.</i>	(TPC-2 $\gamma$ Collab.) JP
ANDO	86	PRL 57 1296	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
ARMSTRONG	84	PL 146B 273	T.A. Armstrong <i>et al.</i>	(ATHU, BARI, BIRM+) JP
BITYUKOV	84	SJNP 39 735	S. Bitjukov <i>et al.</i>	(SERP)
Translated from YAF 39 1165.				
CHAUVAT	84	PL 148B 382	P. Chauvat <i>et al.</i>	(CERN, CLER, UCLA+)
JENNI	83	PR D27 1031	P. Jenni <i>et al.</i>	(SLAC, LBL)
BROMBERG	80	PR D22 1513	C.M. Bromberg <i>et al.</i>	(CIT, FNAL, ILLC+)
DIONISI	80	NP B169 1	C. Dionisi <i>et al.</i>	(CERN, MADR, CDEF+)
CORDEN	78	NP B144 253	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TEL+)
DEFOIX	72	NP B44 125	C. Defoix <i>et al.</i>	(CDEF, CERN)
DAHL	67	PR 163 1377	O.I. Dahl <i>et al.</i>	(LRL) IJP
Also		PRL 14 1074	D.H. Miller <i>et al.</i>	(LRL, UCB)

$$\omega(1420) \quad I^G(J^{PC}) = 0^-(1^--)$$

## $\omega(1420)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>(1400–1450) OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1470 \pm 50$	13.1k	<sup>1</sup> AULCHENKO	15A SND	$1.05\text{--}1.80 e^+ e^- \rightarrow \pi^+\pi^-\pi^0$
$1382 \pm 23 \pm 70$		AUBERT	07Au BABR	$10.6 e^+ e^- \rightarrow \omega\pi^+\pi^-\gamma$
$1350 \pm 20 \pm 20$		AUBERT,B	04N BABR	$10.6 e^+ e^- \rightarrow \pi^+\pi^-\pi^0\gamma$
$1400 \pm 50 \pm 130$	1.2M	<sup>2</sup> ACHASOV	03D RVUE	$0.44\text{--}2.00 e^+ e^- \rightarrow \pi^+\pi^-\pi^0$
$1450 \pm 10$		<sup>3</sup> HENNER	02 RVUE	$1.2\text{--}2.0 e^+ e^- \rightarrow \rho\pi, \omega\pi\pi$
$1373 \pm 70$	177	<sup>4</sup> AKHMETSHIN	00D CMD2	$1.2\text{--}1.38 e^+ e^- \rightarrow \rho\pi$
$1370 \pm 25$	5095	ANISOVICH	00H SPEC	$0.0 \rho\bar{p} \rightarrow \omega\pi^0\pi^0\pi^0$
$1400^{+100}_{-200}$		<sup>5</sup> ACHASOV	98H RVUE	$e^+ e^- \rightarrow \pi^+\pi^-\pi^0$
$\sim 1400$		<sup>6</sup> ACHASOV	98H RVUE	$e^+ e^- \rightarrow \omega\pi^+\pi^-$
$\sim 1460$		<sup>7</sup> ACHASOV	98H RVUE	$e^+ e^- \rightarrow K^+ K^-$
$1440 \pm 70$		<sup>8</sup> CLEGG	94 RVUE	
$1419 \pm 31$	315	<sup>9</sup> ANTONELLI	92 DM2	$1.34\text{--}2.4 e^+ e^- \rightarrow \rho\pi$

<sup>1</sup> From a fit with contributions from  $\omega(782)$ ,  $\phi(1020)$ ,  $\omega(1420)$ , and  $\omega(1650)$ .<sup>2</sup> From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the  $\pi^+\pi^-\pi^0$  and ANTONELLI 92 on the  $\omega\pi^+\pi^-$  final states. Supersedes ACHASOV 99E and ACHASOV 02E.<sup>3</sup> Using results of CORDIER 81 and preliminary data of DOLINSKY 91 and ANTONELLI 92.<sup>4</sup> Using the data of AKHMETSHIN 00D and ANTONELLI 92. The  $\rho\pi$  dominance for the energy dependence of the  $\omega(1420)$  and  $\omega(1650)$  width assumed.<sup>5</sup> Using data from BARKOV 87, DOLINSKY 91, and ANTONELLI 92.<sup>6</sup> Using the data from ANTONELLI 92.<sup>7</sup> Using the data from IVANOV 81 and BISELLO 88b.<sup>8</sup> From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.<sup>9</sup> From a fit to two Breit-Wigner functions interfering between them and with the  $\omega, \phi$  tails with fixed (+, −, +) phases.

# Meson Particle Listings

## $\omega(1420)$ , $f_2(1430)$

$\omega(1420)$ WIDTH					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>(180-250) OUR ESTIMATE</b>					
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
880±170	13.1k	<sup>10</sup> AULCHENKO	15A	SND	$1.05\text{--}1.80\ e^+e^- \rightarrow \frac{\pi^+\pi^-\pi^0}{\pi^+\pi^-\pi^0}$
130± 50±100		AUBERT	07AU	BABR	$10.6\ e^+e^- \rightarrow \omega\pi^+\pi^-\gamma$
450± 70± 70		AUBERT,B	04N	BABR	$10.6\ e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma$
$870^{+500}_{-300}\pm 450$	1.2M	<sup>11</sup> ACHASOV	03D	RVUE	$0.44\text{--}2.00\ e^+e^- \rightarrow \frac{\pi^+\pi^-\pi^0}{\pi^+\pi^-\pi^0}$
199± 15		<sup>12</sup> HENNER	02	RVUE	$1.2\text{--}2.0\ e^+e^- \rightarrow \rho\pi, \omega\pi\pi$
188± 45	177	<sup>13</sup> AKHMETSHIN	00D	CMD2	$1.2\text{--}1.38\ e^+e^- \rightarrow \frac{\omega\pi^+\pi^-}{\omega\pi^+\pi^-}$
$360^{+100}_{-60}$	5095	ANISOVICH	00H	SPEC	$0.0\ p\bar{p} \rightarrow \omega\pi^0\pi^0\pi^0$
240± 70		<sup>14</sup> CLEGG	94	RVUE	
174± 59	315	<sup>15</sup> ANTONELLI	92	DM2	$1.34\text{--}2.4e^+e^- \rightarrow \rho\pi$
<sup>10</sup> From a fit with contributions from $\omega(782)$ , $\phi(1020)$ , $\omega(1420)$ , and $\omega(1650)$ .					
<sup>11</sup> From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the $\pi^+\pi^-\pi^0$ and ANTONELLI 92 on the $\omega\pi^+\pi^-$ final states. Supersedes ACHASOV 99E and ACHASOV 02E.					
<sup>12</sup> Using results of CORDIER 81 and preliminary data of DOLINSKY 91 and ANTONELLI 92.					
<sup>13</sup> Using the data of AKHMETSHIN 00D and ANTONELLI 92. The $\rho\pi$ dominance for the energy dependence of the $\omega(1420)$ and $\omega(1650)$ width assumed.					
<sup>14</sup> From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.					
<sup>15</sup> From a fit to two Breit-Wigner functions interfering between them and with the $\omega,\phi$ tails with fixed (+,-,+) phases.					

$\omega(1420)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\rho\pi$	dominant
$\Gamma_2$ $\omega\pi\pi$	seen
$\Gamma_3$ $b_1(1235)\pi$	seen
$\Gamma_4$ $e^+e^-$	seen
$\Gamma_5$ $\pi^0\gamma$	

$\omega(1420)$ $\Gamma(i)\Gamma(e^+e^-)/\Gamma^2(\text{total})$					
$\Gamma(\rho\pi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma \times \Gamma_4/\Gamma$	VALUE (units $10^{-6}$ )	EVTs	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.73 ±0.08		13.1k	16	AULCHENKO	15A SND 1.05-1.80 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.82 ±0.05 ±0.06				AUBERT,B	04N BABR 10.6 $e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma$
0.65 ±0.13 ±0.21		1.2M	17,18	ACHASOV	03D RVUE 0.44-2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.625±0.160			19,20	CLEGG	94 RVUE
0.466±0.178			21,22	ANTONELLI	92 DM2 1.34-2.4 $e^+e^- \rightarrow \rho\pi$
16 From a fit with contributions from $\omega(782)$ , $\phi(1020)$ , $\omega(1420)$ , and $\omega(1650)$ .					
17 Calculated by us from the cross section at the peak.					
18 From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the $\pi^+\pi^-\pi^0$ and ANTONELLI 92 on the $\omega\pi^+\pi^-$ final states. Supersedes ACHASOV 99E and ACHASOV 02E.					
19 From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.					
20 From the partial and leptonic width given by the authors.					
21 From a fit to two Breit-Wigner functions interfering between them and with the $\omega,\phi$ tails with fixed (+,-,+) phases.					
22 From the product of the leptonic width and partial branching ratio given by the authors.					

$\Gamma(\omega\pi\pi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					
$\Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma \times \Gamma_4/\Gamma$	VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
19.7±5.7				AUBERT	07AU BABR 10.6 $e^+e^- \rightarrow \omega\pi^+\pi^-\gamma$
1.9±1.9			23	AKHMETSHIN	00D CMD2 1.2-2.4 $e^+e^- \rightarrow \omega\pi^+\pi^-$
23 Using the data of AKHMETSHIN 00D and ANTONELLI 92. The $\rho\pi$ dominance for the energy dependence of the $\omega(1420)$ and $\omega(1650)$ width assumed.					

$\Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma_4/\Gamma$
VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$2.03^{+0.70}_{-0.75}$	24	AKHMETSHIN 05	CMD2	0.60-1.38	$e^+e^- \rightarrow \pi^0\gamma$
24 Using 1420 MeV and 220 MeV for the $\omega(1420)$ mass and width.					

$\omega(1420)$ BRANCHING RATIOS	
$\Gamma(\omega\pi\pi)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.301±0.029	25 HENNER 02 RVUE 1.2-2.0 $e^+e^- \rightarrow \rho\pi, \omega\pi\pi$
possibly seen	AKHMETSHIN 00D CMD2 $e^+e^- \rightarrow \omega\pi^+\pi^-$

$\Gamma(\omega\pi\pi)/\Gamma(b_1(1235)\pi)$					$\Gamma_2/\Gamma_3$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.60\pm0.16$	5095	ANISOVICH	00H	SPEC	$0.0\ p\bar{p} \rightarrow \omega\pi^0\pi^0$

$\Gamma(\rho\pi)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.699±0.029	25 HENNER	02	RVUE	1.2-2.0 $e^+e^- \rightarrow \rho\pi, \omega\pi\pi$	

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (units $10^{-7}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$\sim 6.6$	1.2M	<sup>26,27</sup> ACHASOV	03D RVUE	$0.44\text{--}2.00\ e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
$23 \pm 1$		<sup>25</sup> HENNER	02 RVUE	$1.2\text{--}2.0\ e^+e^- \rightarrow \rho\pi, \omega\pi\pi$	
<sup>25</sup> Assuming that the $\omega(1420)$ decays into $\rho\pi$ and $\omega\pi\pi$ only.					
<sup>26</sup> Calculated by us from the cross section at the peak.					
<sup>27</sup> Assuming that the $\omega(1420)$ decays into $\rho\pi$ only.					

$\omega(1420)$ REFERENCES	
AULCHENKO 15A	JETP 121 27 V.M. Aulchenko <i>et al.</i> (SND Collab.)
AUBERT 07AU	Translated from ZETF 148 34. B. Aubert <i>et al.</i> (BABAR Collab.)
AKHMETSHIN 05	PL B605 26 R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
AUBERT,B 04N	PR D70 072004 B. Aubert <i>et al.</i> (BABAR Collab.)
ACHASOV 03D	PR D68 052006 M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
ACHASOV 02E	PR D66 032001 M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
HENNER 02	EPJ C26 3 V.K. Henner <i>et al.</i>
ACHASOV 01E	PR D63 072002 M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
AKHMETSHIN 00D	PL B489 125 R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
ANISOVICH 00H	PL B485 341 A.V. Anisovich <i>et al.</i>
ACHASOV 99E	PL B462 365 M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
ACHASOV 98H	PR D57 4334 N.N. Achasov, A.A. Kozhevnikov
CLEGG 94	ZPHY C62 455 A.B. Clegg, A. Donnachie (LANC, MCHS)
ANTONELLI 92	ZPHY C56 15 A. Antonelli <i>et al.</i> (DM2 Collab.)
DOLINSKY 91	PRPL 202 99 S.I. Dolinsky <i>et al.</i> (NOVO)
BISELLO 88B	ZPHY C39 13 D. Bisello <i>et al.</i> (PADO, CLER, FRAS+)
BARKOV 87	JETPL 46 164 L.M. Barkov <i>et al.</i> (NOVO)
CORDIER 81	PL 106B 155 A. Cordier <i>et al.</i> (ORSAY)
IVANOV 81	PL 107B 297 P.M. Ivanov <i>et al.</i> (NOVO)

<div><math>f_2(1430)</math></div>	$I^G(J^{PC}) = 0^+(2^{++})$
OMITTED FROM SUMMARY TABLE	
This entry lists nearby peaks observed in the $D$ wave of the $K\bar{K}$ and $\pi^+\pi^-$ systems. Needs confirmation.	

$f_2(1430)$ MASS	
VALUE (MeV)	DOCUMENT ID TECN COMMENT
<b>≈ 1430 OUR ESTIMATE</b>	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
1453± 4	1 VLADIMIRSK...01 SPEC 40 $\pi^-p \rightarrow K_S^0 K_S^0 n$
1421± 5	AUGUSTIN 87 DM2 $J/\psi \rightarrow \gamma\pi^+\pi^-$
1480±50	AKESSON 86 SPEC $p\bar{p} \rightarrow p\bar{p}\pi^+\pi^-$
1436 <sup>+26</sup> <sub>-16</sub>	DAUM 84 CNTR 17-18 $\pi^-p \rightarrow K^+K^-n$
1412± 3	DAUM 84 CNTR 63 $\pi^-p \rightarrow K_S^0 K_S^0 n, K^+K^-n$
1439 <sup>+</sup> <sub>- 6</sub>	2 BEUSCH 67 OSPK 5,7,12 $\pi^-p \rightarrow K_S^0 K_S^0 n$
1 $J^{PC} = 0^{++}$ or $2^{++}$ .	
2 Not seen by WETZEL 76.	

$f_2(1430)$ WIDTH	
VALUE (MeV)	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
13± 5	3 VLADIMIRSK...01 SPEC 40 $\pi^-p \rightarrow K_S^0 K_S^0 n$
30± 9	AUGUSTIN 87 DM2 $J/\psi \rightarrow \gamma\pi^+\pi^-$
150±50	AKESSON 86 SPEC $p\bar{p} \rightarrow p\bar{p}\pi^+\pi^-$
81 <sup>+56</sup> <sub>-29</sub>	DAUM 84 CNTR 17-18 $\pi^-p \rightarrow K^+K^-n$
14± 6	DAUM 84 CNTR 63 $\pi^-p \rightarrow K_S^0 K_S^0 n, K^+K^-n$
43 <sup>+17</sup> <sub>-18</sub>	4 BEUSCH 67 OSPK 5,7,12 $\pi^-p \rightarrow K_S^0 K_S^0 n$
3 $J^{PC} = 0^{++}$ or $2^{++}$ .	
4 Not seen by WETZEL 76.	

See key on page 601

## Meson Particle Listings

 $f_2(1430)$ ,  $a_0(1450)$  $f_2(1430)$  DECAY MODES

Mode	
$\Gamma_1$	$K\bar{K}$
$\Gamma_2$	$\pi\pi$

 $f_2(1430)$  REFERENCES

VLADIMIRSKY... 01	PAN 64 1895 Translated from YAF 64 1379.	V.V. Vladimirovsky <i>et al.</i>	
AUGUSTIN 87	ZPHY C36 369	J.E. Augustin <i>et al.</i>	(LALO, CLER, FRAS+)
AKESSON 86	NP B264 154	T. Akesson <i>et al.</i>	(Axial Field Spec. Collab.)
DAUM 84	ZPHY C23 339	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
WETZEL 76	NP B115 208	W. Wetzel <i>et al.</i>	(ETH, CERN, LOIC)
BEUSCH 67	PL 25B 357	W. Beusch <i>et al.</i>	(ETH, CERN)

 $a_0(1450)$ 

$$J^G(J^{PC}) = 1^-(0^{++})$$

See minireview on scalar mesons under  $f_0(500)$ . $a_0(1450)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1474 ±19</b>	<b>OUR AVERAGE</b>			
1480 ±30		ABELE 98	CBAR	0.0 $\bar{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
1470 ±25		<sup>1</sup> AMSLER 95D	CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \pi^0$ , $\pi^0 \eta \eta$ , $\pi^0 \pi^0 \eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1515 ±30		<sup>2</sup> ANISOVICH 09	RVUE	0.0 $\bar{p}p$ , $\pi N$
1316.8 <sup>+</sup> <sub>-1.0</sub> 0.7+24.7 -4.6		<sup>3</sup> UEHARA 09A	BELL	$\gamma\gamma \rightarrow \pi^0 \eta$
1432 ±13 ±25		<sup>4</sup> BUGG 08A	RVUE	$\bar{p}p$
1477 ±10	80k	<sup>5</sup> UMAN 06	E835	5.2 $\bar{p}p \rightarrow \eta \eta \pi^0$
1441 <sup>+</sup> <sub>-15</sub> +40	35280	<sup>2</sup> BAKER 03	SPEC	$\bar{p}p \rightarrow \omega \pi^+ \pi^- \pi^0$
1303 ±16		<sup>6</sup> BARGIOTTI 03	OBLX	$\bar{p}p$
1296 ±10		<sup>7</sup> AMSLER 02	CBAR	0.9 $\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
1565 ±30		<sup>7</sup> ANISOVICH 98B	RVUE	Compilation
1290 ±10		<sup>8</sup> BERTIN 98B	OBLX	0.0 $\bar{p}p \rightarrow K^\pm K_S^0 \pi^\mp$
1450 ±40		AMSLER 94D	CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
1410 ±25		ETKIN 82C	MPS	23 $\pi^- p \rightarrow n 2K_S^0$
~1300		MARTIN 78	SPEC	10 $K^\pm p \rightarrow K_S^0 \pi p$
1255 ±5		<sup>9</sup> CASON 76		

<sup>1</sup> Coupled-channel analysis of AMSLER 95B, AMSLER 95C, and AMSLER 94D.<sup>2</sup> From the pole position.<sup>3</sup> May be a different state.<sup>4</sup> Using data from AMSLER 94D, ABELE 98, and BAKER 03. Supersedes BUGG 94.<sup>5</sup> Statistical error only.<sup>6</sup> Coupled channel analysis of  $\pi^+ \pi^- \pi^0$ ,  $K^+ K^- \pi^0$ , and  $K^\pm K_S^0 \pi^\mp$ .<sup>7</sup> T-matrix pole.<sup>8</sup> Not confirmed by BUGG 08A.<sup>9</sup> Isospin 0 not excluded. $a_0(1450)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>265 ±13</b>	<b>OUR AVERAGE</b>			
265 ±15		ABELE 98	CBAR	0.0 $\bar{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
265 ±30		<sup>10</sup> AMSLER 95D	CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \pi^0$ , $\pi^0 \eta \eta$ , $\pi^0 \pi^0 \eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
230 ±36		<sup>11</sup> ANISOVICH 09	RVUE	0.0 $\bar{p}p$ , $\pi N$
65.0 <sup>+</sup> <sub>-5.4</sub> 2.1+99.1 -32.6		<sup>12</sup> UEHARA 09A	BELL	$\gamma\gamma \rightarrow \pi^0 \eta$
196 ±10 ±10		<sup>13</sup> BUGG 08A	RVUE	$\bar{p}p$
267 ±11	80k	<sup>14</sup> UMAN 06	E835	5.2 $\bar{p}p \rightarrow \eta \eta \pi^0$
110 ±14	35280	<sup>11</sup> BAKER 03	SPEC	$\bar{p}p \rightarrow \omega \pi^+ \pi^- \pi^0$
92 ±16		<sup>15</sup> BARGIOTTI 03	OBLX	$\bar{p}p$
81 ±21		<sup>16</sup> AMSLER 02	CBAR	0.9 $\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
292 ±40		<sup>16</sup> ANISOVICH 98B	RVUE	Compilation
80 ±5		<sup>17</sup> BERTIN 98B	OBLX	0.0 $\bar{p}p \rightarrow K^\pm K_S^0 \pi^\mp$
270 ±40		AMSLER 94D	CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \eta$
230 ±30		ETKIN 82C	MPS	23 $\pi^- p \rightarrow n 2K_S^0$
~250		MARTIN 78	SPEC	10 $K^\pm p \rightarrow K_S^0 \pi p$
79 ±10		<sup>18</sup> CASON 76		

<sup>10</sup> Coupled-channel analysis of AMSLER 95B, AMSLER 95C, and AMSLER 94D.<sup>11</sup> From the pole position.<sup>12</sup> May be a different state.<sup>13</sup> Using data from AMSLER 94D, ABELE 98, and BAKER 03. Supersedes BUGG 94.<sup>14</sup> Statistical error only.<sup>15</sup> Coupled channel analysis of  $\pi^+ \pi^- \pi^0$ ,  $K^+ K^- \pi^0$ , and  $K^\pm K_S^0 \pi^\mp$ .<sup>16</sup> T-matrix pole.<sup>17</sup> Not confirmed by BUGG 08A.<sup>18</sup> Isospin 0 not excluded. $a_0(1450)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$\pi\eta$ seen
$\Gamma_2$	$\pi\eta'(958)$ seen
$\Gamma_3$	$K\bar{K}$ seen
$\Gamma_4$	$\omega\pi\pi$ seen
$\Gamma_5$	$a_0(980)\pi\pi$ seen
$\Gamma_6$	$\gamma\gamma$ seen

 $a_0(1450)$   $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$  $\Gamma(\pi\eta) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  $\Gamma_1\Gamma_6/\Gamma$ 

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
432 ±6 <sup>+</sup> <sub>-256</sub> 1073	<sup>19</sup> UEHARA	09A	BELL $\gamma\gamma \rightarrow \pi^0 \eta$
<sup>19</sup> May be a different state.			

 $a_0(1450)$  BRANCHING RATIOS $\Gamma(\pi\eta'(958))/\Gamma(\pi\eta)$  $\Gamma_2/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.35 ±0.16</b>	<sup>20</sup> ABELE 98	CBAR	0.0 $\bar{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.43 ±0.19	ABELE 97C	CBAR	0.0 $\bar{p}p \rightarrow \pi^0 \pi^0 \eta'$
<sup>20</sup> Using $\pi^0 \eta$ from AMSLER 94D.			

 $\Gamma(K\bar{K})/\Gamma(\pi\eta)$  $\Gamma_3/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.88 ±0.23</b>	<sup>21</sup> ABELE 98	CBAR	0.0 $\bar{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
<sup>21</sup> Using $\pi^0 \eta$ from AMSLER 94D.			

 $\Gamma(\omega\pi\pi)/\Gamma(\pi\eta)$  $\Gamma_4/\Gamma_1$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
10.7 ±2.3	35280	<sup>22</sup> BAKER 03	SPEC	$\bar{p}p \rightarrow \omega \pi^+ \pi^- \pi^0$
<sup>22</sup> Using results on $\bar{p}p \rightarrow a_0(1450)^0 \pi^0$ , $a_0(1450) \rightarrow \eta \pi^0$ from ABELE 96C and assuming the $\omega\rho$ mechanism for the $\omega\pi\pi$ state.				

 $\Gamma(a_0(980)\pi\pi)/\Gamma_{\text{total}}$  $\Gamma_5/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
seen	BUGG 08A	RVUE	$\bar{p}p$

 $\Gamma(a_0(980)\pi\pi)/\Gamma(\pi\eta)$  $\Gamma_5/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
≤ 4.3	ANISOVICH 01	RVUE	0	$\bar{p}p \rightarrow \eta 2\pi^+ 2\pi^-$

 $\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  $\Gamma_6/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
seen	<sup>23</sup> UEHARA	09A	BELL $\gamma\gamma \rightarrow \pi^0 \eta$
<sup>23</sup> May be a different state.			

 $a_0(1450)$  REFERENCES

ANISOVICH 09	IJMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
UEHARA 09A	PR D80 032001	S. Uehara <i>et al.</i>	(BELLE Collab.)
BUGG 08A	PR D78 074023	D.V. Bugg	(LOQM)
UMAN 06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
BAKER 03	PL B563 140	C.A. Baker <i>et al.</i>	
BARGIOTTI 03	EPJ C26 371	M. Bargiotti <i>et al.</i>	(OBELIX Collab.)
AMSLER 02	EPJ C23 29	C. Amisler <i>et al.</i>	
ANISOVICH 01	NP A690 567	A.V. Anisovich <i>et al.</i>	
ABELE 98	PR D57 3860	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ANISOVICH 98B	SPP 41 419	V.V. Anisovich <i>et al.</i>	
Translated from UFN 168 481.			
BERTIN 98B	PL B434 180	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ABELE 97C	PL B404 179	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE 96C	NP A609 562	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER 95B	PL B342 433	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER 95C	PL B353 571	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER 95D	PL B355 425	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER 94D	PL B333 277	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
BUGG 94	PR D50 4412	D.V. Bugg <i>et al.</i>	(LOQM)
ETKIN 82C	PR D25 2446	A. Etkin <i>et al.</i>	(BNL, CUNY, TUFTS, VAND)
MARTIN 78	NP B134 392	A.D. Martin <i>et al.</i>	(DURH, GEVA)
CASON 76	PRL 36 1485	N.M. Cason <i>et al.</i>	(NDAM, ANL)

# Meson Particle Listings

## $\rho(1450)$

**$\rho(1450)$**

See our mini-review under the  $\rho(1700)$ .

### $\rho(1450)$ MASS

**1465 ± 25 OUR ESTIMATE** This is only an educated guess; the error given is larger than the error on the average of the published values.

### $\eta\rho^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1497 ± 14	<sup>1</sup> AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
1421 ± 15	<sup>2</sup> AKHMETSHIN 00D	CMD2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
1470 ± 20	ANTONELLI 88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
1446 ± 10	FUKUI 88	SPEC	$8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

<sup>1</sup> Using the data of AKHMETSHIN 01B on  $e^+e^- \rightarrow \eta\gamma$ , AKHMETSHIN 00D and ANTONELLI 88 on  $e^+e^- \rightarrow \eta\pi^+\pi^-$ .

<sup>2</sup> Using the data of ANTONELLI 88, DOLINSKY 91, and AKHMETSHIN 00D. The energy-independent width of the  $\rho(1450)$  and  $\rho(1700)$  mesons assumed.

### $\omega\pi$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1544 ± 22 +11 -46	821	<sup>1</sup> MATVIENKO 15	BELL	$\overline{B}^0 \rightarrow D^{*+}\omega\pi^-$
1491 ± 19	7815	<sup>2</sup> ACHASOV 13	SND	$1.05^{+2.00}_{-0.00}e^+e^- \rightarrow \pi^0\pi^0\gamma$
1582 ± 17 ± 25	2382	<sup>3</sup> AKHMETSHIN 03B	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
1349 ± 25 +10 -5	341	<sup>4</sup> ALEXANDER 01B	CLE2	$B \rightarrow D^{(*)}\omega\pi^-$
1523 ± 10		<sup>5</sup> EDWARDS 00A	CLE2	$\tau^- \rightarrow \omega\pi^-\nu_\tau$
1463 ± 25		<sup>6</sup> CLEGG 94	RVUE	
1250		<sup>7</sup> ASTON 80c	OMEG	$20\text{--}70\gamma p \rightarrow \omega\pi^0 p$
1290 ± 40		<sup>7</sup> BARBER 80c	SPEC	$3\text{--}5\gamma p \rightarrow \omega\pi^0 p$

<sup>1</sup> Using Breit-Wigner parameterization of the  $\rho(1450)$  and assuming equal probabilities of the  $\rho(1450) \rightarrow \pi\pi$  and  $\rho(1450) \rightarrow \omega\pi$  decays.

<sup>2</sup> From a phenomenological model based on vector meson dominance with the interfering  $\rho(1450)$  and  $\rho(1700)$  and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.

<sup>3</sup> Using the data of AKHMETSHIN 03B and BISELLO 91B assuming the  $\omega\pi^0$  and  $\pi^+\pi^-$  mass dependence of the total width.  $\rho(1700)$  mass and width fixed at 1700 MeV and 240 MeV, respectively.

<sup>4</sup> Using Breit-Wigner parameterization of the  $\rho(1450)$  and assuming the  $\omega\pi^-$  mass dependence for the total width.

<sup>5</sup> Mass-independent width parameterization.  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.

<sup>6</sup> Using data from BISELLO 91B, DOLINSKY 86 and ALBRECHT 87L.

<sup>7</sup> Not separated from  $b_1(1235)$ , not pure  $J^P = 1^-$  effect.

### $4\pi$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1435 ± 40	ABELE 01B	CBAR	$0.0\overline{p}n \rightarrow 2\pi^-2\pi^0\pi^+$
1350 ± 50	ACHASOV 97	RVUE	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1449 ± 4	<sup>1</sup> ARMSTRONG 89E	OMEG	$300pp \rightarrow p\rho 2(\pi^+\pi^-)$

<sup>1</sup> Not clear whether this observation has  $l=1$  or 0.

### $\pi\pi$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1350 ± 20 +20 -30	63.5k	<sup>1</sup> ABRAMOWICZ12	ZEUS	$e p \rightarrow e\pi^+\pi^-\rho$
1493 ± 15		<sup>2</sup> LEES 12G	BABR	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
1446 ± 7 ± 28	5.4M	<sup>3,4</sup> FUJIKAWA 08	BELL	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
1328 ± 15		<sup>5</sup> SCHAEEL 05c	ALEP	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
1406 ± 15	87k	<sup>3,6</sup> ANDERSON 00A	CLE2	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
~ 1368		<sup>7</sup> ABELE 99c	CBAR	$0.0\overline{p}d \rightarrow \pi^+\pi^-\pi^-p$
1348 ± 33		BERTIN 98	OBLX	$0.05\text{--}0.405\overline{p}p \rightarrow 2\pi^+\pi^-$
1411 ± 14		<sup>8</sup> ABELE 97	CBAR	$\overline{p}n \rightarrow \pi^-\pi^0\pi^0$
1370 ± 90 -70		ACHASOV 97	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
1359 ± 40		<sup>6</sup> BERTIN 97c	OBLX	$0.0\overline{p}p \rightarrow \pi^+\pi^-\pi^0$
1282 ± 37		BERTIN 97D	OBLX	$0.05\overline{p}p \rightarrow 2\pi^+2\pi^-$
1424 ± 25		BISELLO 89	DM2	$e^+e^- \rightarrow \pi^+\pi^-$
1265.5 ± 75.3		DUBNICKA 89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
1292 ± 17		<sup>9</sup> KURDADZE 83	OLYA	$0.64\text{--}1.4e^+e^- \rightarrow \pi^+\pi^-$

<sup>1</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho\text{--}\omega$  interference.

<sup>2</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>3</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>4</sup>  $|F_\pi(0)|^2$  fixed to 1.

<sup>5</sup> From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEEL 05c and  $e^+e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05.  $\rho(1700)$  mass and width fixed at 1713 MeV and 235 MeV, respectively. Supersedes BARATE 97M.

<sup>6</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV, respectively.

<sup>7</sup>  $\rho(1700)$  mass and width fixed at 1780 MeV and 275 MeV respectively.

<sup>8</sup> T-matrix pole.

<sup>9</sup> Using for  $\rho(1700)$  mass and width  $1600 \pm 20$  and  $300 \pm 10$  MeV respectively.

### $K\overline{K}$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1422.8 ± 6.5	27k	<sup>1</sup> ABELE 99D	CBAR	0.0	$\overline{p}p \rightarrow K^+K^-\pi^0$
<sup>1</sup> K-matrix pole. Isospin not determined, could be $\omega(1420)$ .					

### $K\overline{K}^*(892) + \text{c.c.}$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1505 ± 19 ± 7	AUBERT 08s	BABR	$10.6e^+e^- \rightarrow K\overline{K}^*(892)\gamma$

### $\rho(1450)$ WIDTH

**400 ± 60 OUR ESTIMATE** This is only an educated guess; the error given is larger than the error on the average of the published values.

### $\eta\rho^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
226 ± 44	<sup>1</sup> AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
211 ± 31	<sup>2</sup> AKHMETSHIN 00D	CMD2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
230 ± 30	ANTONELLI 88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
60 ± 15	FUKUI 88	SPEC	$8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

<sup>1</sup> Using the data of AKHMETSHIN 01B on  $e^+e^- \rightarrow \eta\gamma$ , AKHMETSHIN 00D and ANTONELLI 88 on  $e^+e^- \rightarrow \eta\pi^+\pi^-$ .

<sup>2</sup> Using the data of ANTONELLI 88, DOLINSKY 91, and AKHMETSHIN 00D. The energy-independent width of the  $\rho(1450)$  and  $\rho(1700)$  mesons assumed.

### $\omega\pi$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
303 ± 31 +69 -52 -7	821	<sup>1</sup> MATVIENKO 15	BELL	$\overline{B}^0 \rightarrow D^{*+}\omega\pi^-$
429 ± 42 ± 10	2382	<sup>2</sup> AKHMETSHIN 03B	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
547 ± 86 +46 -45	341	<sup>3</sup> ALEXANDER 01B	CLE2	$B \rightarrow D^{(*)}\omega\pi^-$
400 ± 35		<sup>4</sup> EDWARDS 00A	CLE2	$\tau^- \rightarrow \omega\pi^-\nu_\tau$
311 ± 62		<sup>5</sup> CLEGG 94	RVUE	
300		<sup>6</sup> ASTON 80c	OMEG	$20\text{--}70\gamma p \rightarrow \omega\pi^0 p$
320 ± 100		<sup>6</sup> BARBER 80c	SPEC	$3\text{--}5\gamma p \rightarrow \omega\pi^0 p$

<sup>1</sup> Using Breit-Wigner parameterization of the  $\rho(1450)$  and assuming equal probabilities of the  $\rho(1450) \rightarrow \pi\pi$  and  $\rho(1450) \rightarrow \omega\pi$  decays.

<sup>2</sup> Using the data of AKHMETSHIN 03B and BISELLO 91B assuming the  $\omega\pi^0$  and  $\pi^+\pi^-$  mass dependence of the total width.  $\rho(1700)$  mass and width fixed at 1700 MeV and 240 MeV, respectively.

<sup>3</sup> Using Breit-Wigner parameterization of the  $\rho(1450)$  and assuming the  $\omega\pi^-$  mass dependence for the total width.

<sup>4</sup> Mass-independent width parameterization.  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.

<sup>5</sup> Using data from BISELLO 91B, DOLINSKY 86 and ALBRECHT 87L.

<sup>6</sup> Not separated from  $b_1(1235)$ , not pure  $J^P = 1^-$  effect.

### $4\pi$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
325 ± 100	ABELE 01B	CBAR	$0.0\overline{p}n \rightarrow 2\pi^-2\pi^0\pi^+$

### $\pi\pi$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
460 ± 30 +40 -45	63.5k	<sup>1</sup> ABRAMOWICZ12	ZEUS	$e p \rightarrow e\pi^+\pi^-\rho$
427 ± 31		<sup>2</sup> LEES 12G	BABR	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
434 ± 16 ± 60	5.4M	<sup>3,4</sup> FUJIKAWA 08	BELL	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
468 ± 41		<sup>5</sup> SCHAEEL 05c	ALEP	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
455 ± 41	87k	<sup>3,6</sup> ANDERSON 00A	CLE2	$\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
~ 374		<sup>7</sup> ABELE 99c	CBAR	$0.0\overline{p}d \rightarrow \pi^+\pi^-\pi^-p$
275 ± 10		BERTIN 98	OBLX	$0.05\text{--}0.405\overline{p}p \rightarrow \pi^+\pi^-\pi^-$
343 ± 20		<sup>8</sup> ABELE 97	CBAR	$\overline{p}n \rightarrow \pi^-\pi^0\pi^0$
310 ± 40		<sup>6</sup> BERTIN 97c	OBLX	$0.0\overline{p}p \rightarrow \pi^+\pi^-\pi^0$
236 ± 36		BERTIN 97D	OBLX	$0.05\overline{p}p \rightarrow 2\pi^+2\pi^-$
269 ± 31		BISELLO 89	DM2	$e^+e^- \rightarrow \pi^+\pi^-$
391 ± 70		DUBNICKA 89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
218 ± 46		<sup>9</sup> KURDADZE 83	OLYA	$0.64\text{--}1.4e^+e^- \rightarrow \pi^+\pi^-$

<sup>1</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho\text{--}\omega$  interference.

<sup>2</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>3</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>4</sup>  $|F_\pi(0)|^2$  fixed to 1.

<sup>5</sup> From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEEL 05c and  $e^+e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05.  $\rho(1700)$  mass and width fixed at 1713 MeV and 235 MeV, respectively. Supersedes BARATE 97M.

<sup>6</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV, respectively.

<sup>7</sup>  $\rho(1700)$  mass and width fixed at 1780 MeV and 275 MeV respectively.

<sup>8</sup> T-matrix pole.

<sup>9</sup> Using for  $\rho(1700)$  mass and width  $1600 \pm 20$  and  $300 \pm 10$  MeV respectively.

See key on page 601

Meson Particle Listings

$\rho(1450)$

$K\overline{K}$ MODE					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$146.5 \pm 10.5$	27k	<sup>1</sup> ABELE	99D	CBAR	$\pm$ 0.0 $\overline{p}p \rightarrow K^+ K^- \pi^0$
<sup>1</sup> K-matrix pole. Isospin not determined, could be $\omega(1420)$ .					

$K\overline{K}^*(892) + \text{c.c. MODE}$				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$418 \pm 25 \pm 4$	AUBERT	08s	BABR	$10.6 \text{ e}^+ \text{e}^- \rightarrow K\overline{K}^*(892)\gamma$

$\rho(1450)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \pi\pi$	seen	
$\Gamma_2 4\pi$	seen	
$\Gamma_3 \omega\pi$		
$\Gamma_4 a_1(1260)\pi$		
$\Gamma_5 h_1(1170)\pi$		
$\Gamma_6 \pi(1300)\pi$		
$\Gamma_7 \rho\rho$		
$\Gamma_8 \rho(\pi\pi) \text{ s-wave}$		
$\Gamma_9 e^+ e^-$	seen	
$\Gamma_{10} \eta\rho$	seen	
$\Gamma_{11} a_2(1320)\pi$	not seen	
$\Gamma_{12} K\overline{K}$	not seen	
$\Gamma_{13} K\overline{K}^*(892) + \text{c.c.}$	possibly seen	
$\Gamma_{14} \eta\gamma$	seen	
$\Gamma_{15} f_0(500)\gamma$	not seen	
$\Gamma_{16} f_0(980)\gamma$	not seen	
$\Gamma_{17} f_0(1370)\gamma$	not seen	
$\Gamma_{18} f_2(1270)\gamma$	not seen	

$\rho(1450) \Gamma(i)\Gamma(e^+ e^-)/\Gamma(\text{total})$		
$\Gamma(\pi\pi) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	$\Gamma_1\Gamma_9/\Gamma$	
VALUE (keV)	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
0.12	<sup>1</sup> DIEKMANN	88 RVUE $e^+ e^- \rightarrow \pi^+ \pi^-$
$0.027^{+0.015}_{-0.010}$	<sup>2</sup> KURDADZE	83 OLYA 0.64–1.4 $e^+ e^- \rightarrow \pi^+ \pi^-$

$\Gamma(\eta\rho) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (eV)	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$74 \pm 20$	<sup>3</sup> AKHMETSHIN 00D	CMD2 $e^+ e^- \rightarrow \eta\pi^+ \pi^-$
$91 \pm 19$	ANTONELLI 88	DM2 $e^+ e^- \rightarrow \eta\pi^+ \pi^-$

$\Gamma(\eta\gamma) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (eV)	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<16.4	<sup>4</sup> AKHMETSHIN 05	CMD2 0.60–1.38 $e^+ e^- \rightarrow \eta\gamma$
$2.2 \pm 0.5 \pm 0.3$	<sup>5</sup> AKHMETSHIN 01B	CMD2 $e^+ e^- \rightarrow \eta\gamma$

$\Gamma(K\overline{K}^*(892) + \text{c.c.}) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (eV)	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$127 \pm 15 \pm 6$	AUBERT	08s BABR 10.6 $e^+ e^- \rightarrow K^+ \overline{K}^*(892) \gamma$
<sup>1</sup> Using total width = 235 MeV.		
<sup>2</sup> Using for $\rho(1700)$ mass and width $1600 \pm 20$ and $300 \pm 10$ MeV respectively.		
<sup>3</sup> Using the data of ANTONELLI 88, DOLINSKY 91, and AKHMETSHIN 00D. The energy-independent width of the $\rho(1450)$ and $\rho(1700)$ mesons assumed.		
<sup>4</sup> From $2\gamma$ decay mode of $\eta$ using 1465 MeV and 310 MeV for the $\rho(1450)$ mass and width. Recalculated by us.		
<sup>5</sup> Using the data of AKHMETSHIN 01B on $e^+ e^- \rightarrow \eta\gamma$ , AKHMETSHIN 00D and ANTONELLI 88 on $e^+ e^- \rightarrow \eta\pi^+ \pi^-$ . Recalculated by us using width of 226 MeV.		

$\rho(1450) \Gamma(i)/\Gamma(\text{total}) \times \Gamma(e^+ e^-)/\Gamma(\text{total})$		
$\Gamma(\omega\pi)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma \times \Gamma_9/\Gamma$	
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$5.3 \pm 0.4$	7815	<sup>1</sup> ACHASOV 13 SND 1.05–2.00 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
$\Gamma(\eta\rho)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma \times \Gamma_9/\Gamma$	
VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$4.3^{+1.1}_{-0.9} \pm 0.2$	4.9k	<sup>2</sup> AULCHENKO 15 SND 1.22–2.00 $e^+ e^- \rightarrow \eta\pi^+ \pi^-$

$\Gamma(f_0(500)\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID TECN COMMENT
<4.0	90	ACHASOV 11 SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$

$\Gamma(f_0(980)\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID TECN COMMENT
<2.6	90	ACHASOV 11 SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$

$\Gamma(f_0(1370)\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID TECN COMMENT
<3.5	90	ACHASOV 11 SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$

$\Gamma(f_2(1270)\gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID TECN COMMENT
<0.8	90	<sup>3</sup> ACHASOV 11 SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$

<sup>1</sup> From a phenomenological model based on vector meson dominance with the interfering  $\rho(1450)$  and  $\rho(1700)$  and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.

<sup>2</sup> From a fit to the  $e^+ e^- \rightarrow \eta\pi^+ \pi^-$  cross section with vector meson dominance model including  $\rho(770)$ ,  $\rho(1450)$ , and  $\rho(1700)$  decaying exclusively via  $\eta\rho(770)$ . Masses and widths of vector states are fixed to PDG 14. Coupling constants are assumed to be real.

<sup>3</sup> Using Breit-Wigner parametrization of the  $\rho(1450)$  with mass and width of 1465 MeV and 400 MeV, respectively.

$\rho(1450)$ BRANCHING RATIOS		
$\Gamma(\pi\pi)/\Gamma(4\pi)$	$\Gamma_1/\Gamma_2$	
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.37 \pm 0.10$	<sup>1,2</sup> ABELE	01B CBAR 0.0 $\overline{p}n \rightarrow 5\pi$

$\Gamma(\omega\pi)/\Gamma_{\text{total}}$		
VALUE	EVTS	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
seen	821	<sup>3</sup> MATVIENKO 15 BELL $\overline{B}^0 \rightarrow D^{*+} \omega \pi^-$
seen	1.6k	ACHASOV 12 SND $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
$\sim 0.21$		CLEGG 94 RVUE

$\Gamma(\pi\pi)/\Gamma(\omega\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$\sim 0.32$	CLEGG	94 RVUE

$\Gamma(\omega\pi)/\Gamma(4\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<0.14	CLEGG	88 RVUE

$\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.27 \pm 0.08$	<sup>1</sup> ABELE	01B CBAR 0.0 $\overline{p}n \rightarrow 5\pi$

$\Gamma(h_1(1170)\pi)/\Gamma(4\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.08 \pm 0.04$	<sup>1</sup> ABELE	01B CBAR 0.0 $\overline{p}n \rightarrow 5\pi$

$\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.37 \pm 0.13$	<sup>1</sup> ABELE	01B CBAR 0.0 $\overline{p}n \rightarrow 5\pi$

$\Gamma(\rho\rho)/\Gamma(4\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.11 \pm 0.05$	<sup>1</sup> ABELE	01B CBAR 0.0 $\overline{p}n \rightarrow 5\pi$

$\Gamma(\rho(\pi\pi) \text{ s-wave})/\Gamma(4\pi)$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.17 \pm 0.09$	<sup>1</sup> ABELE	01B CBAR 0.0 $\overline{p}n \rightarrow 5\pi$

$\Gamma(\eta\rho)/\Gamma_{\text{total}}$		
VALUE	EVTS	DOCUMENT ID TECN COMMENT
seen	35	<sup>4</sup> ACHASOV 14 SND 1.15–2.00 $e^+ e^- \rightarrow \eta\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<0.04		DONNACHIE 87B RVUE

# Meson Particle Listings

## $\rho(1450)$ , $\eta(1475)$

### $\Gamma(\eta\rho)/\Gamma(\omega\pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.081 \pm 0.020$	<sup>5,6</sup> AULCHENKO 15	SND	$1.22\text{--}2.00\ e^+e^- \rightarrow \eta\pi^+\pi^-$
$\sim 0.24$	<sup>7</sup> DONNACHIE 91	RVUE	
$>2$	FUKUI 91	SPEC	$8.95\ \pi^-p \rightarrow \omega\pi^0n$

### $\Gamma(\pi\pi)/\Gamma(\eta\rho)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.3 \pm 0.4$	<sup>5</sup> AULCHENKO 15	SND	$1.22\text{--}2.00\ e^+e^- \rightarrow \eta\pi^+\pi^-$

### $\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	AMELIN 00	VES	$37\ \pi^-p \rightarrow \eta\pi^+\pi^-n$

### $\Gamma(K\overline{K})/\Gamma(\omega\pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$<0.08$	<sup>7</sup> DONNACHIE 91	RVUE	

### $\Gamma(K\overline{K}^*(892) + \text{c.c.})/\Gamma_{\text{total}}$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
possibly seen	COAN 04	CLEO	$\tau^- \rightarrow K^-\pi^-K^+\nu_\tau$

### $\Gamma(\eta\gamma)/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	35	<sup>4</sup> ACHASOV 14	SND	$1.15\text{--}2.00\ e^+e^- \rightarrow \eta\gamma$

- <sup>1</sup>  $\omega\pi$  not included.  
<sup>2</sup> Using ABELE 97.  
<sup>3</sup> Using Breit-Wigner parameterization of the  $\rho(1450)$  and assuming equal probabilities of the  $\rho(1450) \rightarrow \pi\pi$  and  $\rho(1450) \rightarrow \omega\pi$  decays.  
<sup>4</sup> From a phenomenological model based on vector meson dominance with  $\rho(1450)$  and  $\phi(1680)$  masses and widths from the PDG 12.  
<sup>5</sup> From a fit to the  $e^+e^- \rightarrow \eta\pi^+\pi^-$  cross section with vector meson dominance model including  $\rho(770)$ ,  $\rho(1450)$ , and  $\rho(1700)$  decaying exclusively via  $\eta\rho(770)$ . Masses and widths of vector states are fixed to PDG 14. Coupling constants are assumed to be real.  
<sup>6</sup> Reports the inverse of the quoted value as  $12.3 \pm 3.1$ .  
<sup>7</sup> Using data from BISELLO 91B, DOLINSKY 86 and ALBRECHT 87L.

### $\rho(1450)$ REFERENCES

AULCHENKO 15	PR D91 052013	V.M. Aulchenko <i>et al.</i>	(SND Collab.)
MATVIENKO 15	PR D92 012013	D. Matvienko <i>et al.</i>	(BELLE Collab.)
ACHASOV 14	PR D90 032002	M.N. Achasov <i>et al.</i>	(SND Collab.)
PDG 14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
ACHASOV 13	PR D88 054013	M.N. Achasov <i>et al.</i>	(SND Collab.)
ABRAMOWICZ 12	EPJ C72 1869	H. Abramowicz <i>et al.</i>	(ZEUS Collab.)
ACHASOV 12	JETPL 94 734	M.N. Achasov <i>et al.</i>	
	Translated from ZETFP 94 796		
LEES 12G	PR D86 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
PDG 12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
ACHASOV 11	JETP 113 75	M.N. Achasov <i>et al.</i>	(SND Collab.)
	Translated from ZETFP 140 87.		
AUBERT 08S	PR D77 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
FUJIKAWA 08	PR D78 072006	M. Fujikawa <i>et al.</i>	(BELLE Collab.)
AKHMETSHIN 05	PL B605 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ALOISIO 05	PL B606 12	A. Aloisio <i>et al.</i>	(KLOE Collab.)
SCHAEF 05C	PRPL 421 191	S. Schaefer <i>et al.</i>	(ALEPH Collab.)
AKHMETSHIN 04	PL B578 285	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
COAN 04	PRL 92 232001	T.E. Coan <i>et al.</i>	(CLEO Collab.)
AKHMETSHIN 03B	PL B562 173	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ABELE 01B	EPJ C21 261	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
AKHMETSHIN 01B	PL B509 217	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ALEXANDER 01B	PR D64 092001	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
AKHMETSHIN 00D	PL B489 125	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AMELIN 00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
ANDERSON 00A	PR D61 112002	S. Anderson <i>et al.</i>	(CLEO Collab.)
EDWARDS 00A	PR D61 072003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ABELE 99C	PL B450 275	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE 99D	PL B468 178	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BERTIN 98	PR D57 55	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ABELE 97	PL B391 191	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV 97	PR D55 2663	M.N. Achasov <i>et al.</i>	(NOVM Collab.)
BARATE 97M	ZPHY C76 15	R. Barate <i>et al.</i>	(ALEPH Collab.)
BERTIN 97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
BERTIN 97D	PL B414 220	A. Bertin <i>et al.</i>	(OBELIX Collab.)
CLEGG 94	ZPHY C62 455	A.B. Clegg, A. Donnachie	(LANC, MCHS)
BISELLO 91B	NPBPS B21 111	D. Bisello	(DM2 Collab.)
DOLINSKY 91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
DONNACHIE 91	ZPHY C51 689	A. Donnachie, A.B. Clegg	(MCHS, LANC)
FUKUI 91	PL B257 241	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)

### $\Gamma_{10}/\Gamma_3$

KUHN 90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ARMSTRONG 89E	PL B228 536	T.A. Armstrong, M. Benayoun	(ATHU, BARI, BIRM+)
BISELLO 89	PL B220 321	D. Bisello <i>et al.</i>	(DM2 Collab.)
DUBNICKA 89	JP G15 1349	S. Dubnicka <i>et al.</i>	(JINR, SLOV)
ANTONELLI 88	PL B212 133	A. Antonelli <i>et al.</i>	(DM2 Collab.)
CLEGG 88	ZPHY C40 313	A.B. Clegg, A. Donnachie	(MCHS, LANC)
DIEKMAN 88	PRPL 159 99	B. Diekmann	(BONN)
FUKUI 88	PL B202 441	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
ALBRECHT 87L	PL B185 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DONNACHIE 87B	ZPHY C34 257	A. Donnachie, A.B. Clegg	(MCHS, LANC)
DOLINSKY 86	PL B174 453	S.I. Dolinsky <i>et al.</i>	(NOVO)
BARKOV 85	NP B256 365	L.M. Barkov <i>et al.</i>	(NOVO)
KURDADZE 83	JETPL 37 733	L.M. Kurdadze <i>et al.</i>	(NOVO)
	Translated from ZETFP 37 613.		
ASTON 80C	PL 92B 211	D. Aston	(BONN, CERN, EPOL, GLAS, LANC+)
BARBER 80C	ZPHY C4 169	D.P. Barber <i>et al.</i>	(DARE, LANC, SHEF)
GOUNARIS 68	PRL 21 244	G.J. Gounaris, J.J. Sakurai	

## $\eta(1475)$

$$I^G(J^{PC}) = 0^+(0^{-+})$$

See also the  $\eta(1405)$ .

### $\eta(1475)$ MASS

#### $K\overline{K}\pi$ MODE ( $K^*(892)$ $K$ dominant)

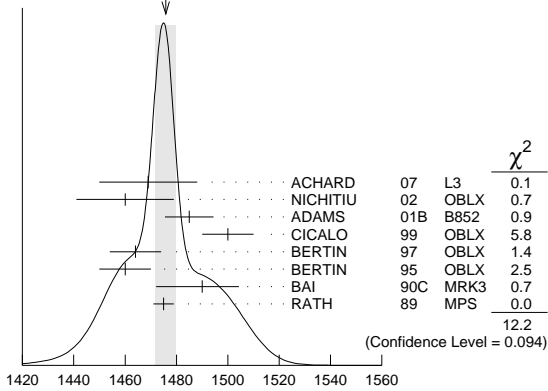
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1476 \pm 4</math> OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
$1469 \pm 14 \pm 13$	74	ACHARD 07	L3	$183\text{--}209\ e^+e^- \rightarrow e^+e^-K_S^0K^\pm\pi^\mp$
$1460 \pm 19$	3651	NICHITIU 02	OBLX	
$1485 \pm 8 \pm 5$	20k	ADAMS 01B	B852	$18\ \text{GeV}\ \pi^-p \rightarrow K^+K^-\pi^0n$
$1500 \pm 10$		CICALO 99	OBLX	$0\ \overline{p}p \rightarrow K^\pm K_S^0\pi^\mp\pi^+\pi^-$
$1464 \pm 10$		BERTIN 97	OBLX	$0\ \overline{p}p \rightarrow K^\pm(K^0)\pi^\mp\pi^+\pi^-$
$1460 \pm 10$		BERTIN 95	OBLX	$0\ \overline{p}p \rightarrow K\overline{K}\pi\pi\pi$
$1490^{+14}_{-8} + \frac{3}{-16}$	1100	BAI 90C	MRK3	$J/\psi \rightarrow \gamma K_S^0K^\pm\pi^\mp$
$1475 \pm 4$		RATH 89	MPS	$21.4\ \pi^-p \rightarrow nK_S^0K^\pm\pi^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1565 \pm 8^{+0}_{-63}$	<sup>1</sup> ABLIKIM 15T	BES3	$J/\psi \rightarrow \gamma K_S^0K_S^0\eta$
$1421 \pm 14$	AUGUSTIN 92	DM2	$J/\psi \rightarrow \gamma K\overline{K}\pi$

<sup>1</sup> Could also be the  $\eta(1405)$ .

WEIGHTED AVERAGE  
 $1476 \pm 4$  (Error scaled by 1.3)



$\eta(1475)$  mass,  $K\overline{K}\pi$  mode ( $K^*(892)$   $K$  dominant) (MeV)

### $\eta(1475)$ WIDTH

#### $K\overline{K}\pi$ MODE ( $K^*(892)$ $K$ dominant)

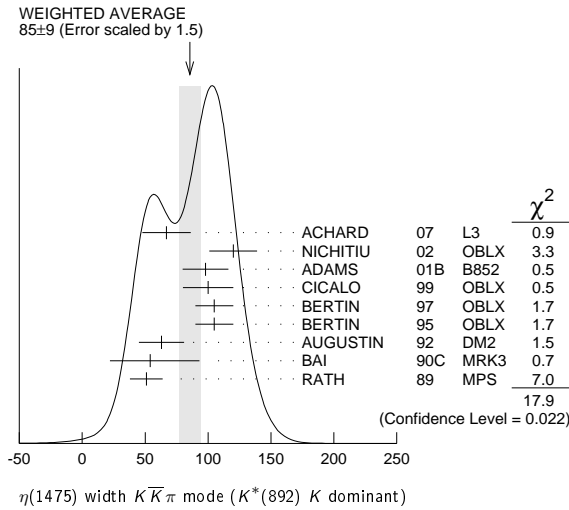
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>85 \pm 9</math> OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
$67 \pm 18 \pm 7$	74	ACHARD 07	L3	$183\text{--}209\ e^+e^- \rightarrow e^+e^-K_S^0K^\pm\pi^\mp$
$120 \pm 19$	3651	NICHITIU 02	OBLX	
$98 \pm 18 \pm 3$	20k	ADAMS 01B	B852	$18\ \text{GeV}\ \pi^-p \rightarrow K^+K^-\pi^0n$
$100 \pm 20$		CICALO 99	OBLX	$0\ \overline{p}p \rightarrow K^\pm K_S^0\pi^\mp\pi^+\pi^-$
$105 \pm 15$		BERTIN 97	OBLX	$0.0\ \overline{p}p \rightarrow K^\pm(K^0)\pi^\mp\pi^+\pi^-$
$105 \pm 15$		BERTIN 95	OBLX	$0\ \overline{p}p \rightarrow K\overline{K}\pi\pi\pi$
$63 \pm 18$		AUGUSTIN 92	DM2	$J/\psi \rightarrow \gamma K\overline{K}\pi$
$54^{+37}_{-21} + \frac{13}{-24}$		BAI 90C	MRK3	$J/\psi \rightarrow \gamma K_S^0K^\pm\pi^\mp$
$51 \pm 13$		RATH 89	MPS	$21.4\ \pi^-p \rightarrow nK_S^0K_S^0\pi^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$54^{+14}_{-13} + \frac{21}{-28}$	<sup>1</sup> ABLIKIM 15T	BES3	$J/\psi \rightarrow \gamma K_S^0K_S^0\eta$
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See key on page 601

## Meson Particle Listings

 $\eta(1475)$ ,  $f_0(1500)$ <sup>1</sup> Could also be the  $\eta(1405)$ . $\eta(1475)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\bar{K}\pi$	dominant
$\Gamma_2$ $K\bar{K}^*(892) + c.c.$	seen
$\Gamma_3$ $a_0(980)\pi$	seen
$\Gamma_4$ $\gamma\gamma$	seen
$\Gamma_5$ $K_S^0 K_S^0 \eta$	possibly seen

 $\eta(1475)$   $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

VALUE (keV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$0.23 \pm 0.05 \pm 0.05$		74	1 ACHARD	07 L3	$183-209 e^+ e^- \rightarrow e^+ e^- K_S^0 K^\pm \pi^\mp$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.089	90	2,3 AHOHE	05 CLE2	10.6 $e^+ e^- \rightarrow e^+ e^- K_S^0 K^\pm \pi^\mp$
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<sup>1</sup> Supersedes ACCIARRI 01G. Compatible with  $K^* K$  decay. Using  $B(K_S^0 \rightarrow \pi^+ \pi^-) = 0.6895$ .<sup>2</sup> Using  $\eta(1475)$  mass of 1481 MeV and width of 48 MeV. The upper limit increases to 0.140 keV if the world average value, 87 MeV, of the width is used.<sup>3</sup> Assuming three-body phase-space decay to  $K_S^0 K^\pm \pi^\mp$ . $\eta(1475)$  BRANCHING RATIOS $\Gamma(K\bar{K}^*(892) + c.c.)/\Gamma(K\bar{K}\pi)$   $\Gamma_2/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.50 \pm 0.10$	1 BAILLON	67 HBC	$0.0 \bar{p} p \rightarrow K\bar{K}\pi\pi$

<sup>1</sup> Data could also refer to  $\eta(1405)$ . $\Gamma(K\bar{K}^*(892) + c.c.)/[\Gamma(K\bar{K}^*(892) + c.c.) + \Gamma(a_0(980)\pi)]$   $\Gamma_2/(\Gamma_2 + \Gamma_3)$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< 0.25	90	EDWARDS	82E CBAL	$J/\psi \rightarrow K^+ K^- \pi^0 \gamma$

 $\eta(1475)$  REFERENCES

ABLIKIM	15T	PRL 115 091803	M. Ablikim <i>et al.</i>	(BES III Collab.)
ACHARD	07	JHEP 0703 018	P. Achard <i>et al.</i>	(L3 Collab.)
AHOHE	05	PR D71 072001	R. Ahohe <i>et al.</i>	(CLEO Collab.)
NICHITIU	02	PL B545 261	F. Nichitiu <i>et al.</i>	(OBLIX Collab.)
ACCIARRI	01G	PL B501 1	M. Acciarri <i>et al.</i>	(L3 Collab.)
ADAMS	01B	PL B516 264	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)
CICALO	99	PL B462 453	C. Cicalo <i>et al.</i>	(OBLIX Collab.)
BERTIN	97	PL B400 226	A. Bertin <i>et al.</i>	(OBLIX Collab.)
BERTIN	95	PL B361 187	A. Bertin <i>et al.</i>	(OBLIX Collab.)
AUGUSTIN	92	PR D46 1951	J.E. Augustin, G. Cosme	(DM2 Collab.)
BAI	90C	PRL 65 2507	Z. Bai <i>et al.</i>	(Mark III Collab.)
RATH	89	PR D40 693	M.G. Rath <i>et al.</i>	(NDAM, BRAN, BNL, CUNY+)
EDWARDS	82E	PRL 49 259	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
BAILLON	67	NC 50A 393	P.H. Baillon <i>et al.</i>	(CERN, CDEF, IRAD)

 $f_0(1500)$ 

$$J^G(J^{PC}) = 0^+(0^{++})$$

See also the mini-reviews on scalar mesons under  $f_0(500)$  (see the index for the page number) and on non- $q\bar{q}$  candidates in PDG 06, Journal of Physics **G33** 1 (2006). $f_0(1500)$  MASS

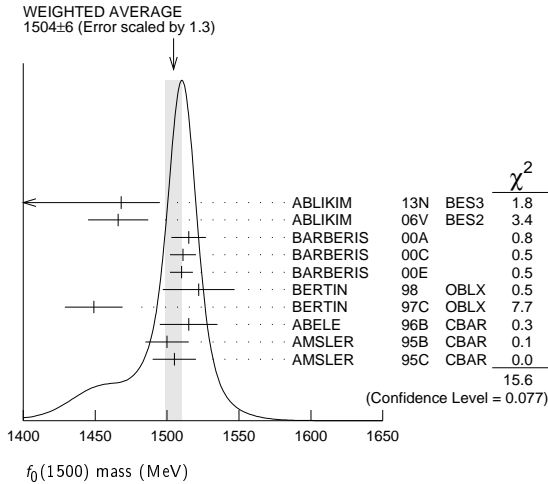
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1504 ± 6 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
$1468^{+14}_{-15} +^{23}_{-74}$	5.5k	1 ABLIKIM	13N BES3	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma \eta \eta$
$1466 \pm 6 \pm 20$		ABLIKIM	06V BES2	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma \pi^+ \pi^-$
$1515 \pm 12$		2 BARBERIS	00A	$450 p p \rightarrow p_f \eta \eta p_S$
$1511 \pm 9$		2,3 BARBERIS	00C	$450 p p \rightarrow p_f 4 \pi p_S$
$1510 \pm 8$		2 BARBERIS	00E	$450 p p \rightarrow p_f \eta \eta p_S$
$1522 \pm 25$		BERTIN	98 OBLX	$0.05-0.405 \pi p \rightarrow \pi^+ \pi^+ \pi^-$
$1449 \pm 20$		2 BERTIN	97C OBLX	$0.0 \bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
$1515 \pm 20$		ABELE	96B CBAR	$0.0 \bar{p} p \rightarrow \pi^0 K_L^0 K_L^0$
$1500 \pm 15$		4 AMSLER	95B CBAR	$0.0 \bar{p} p \rightarrow 3 \pi^0$
$1505 \pm 15$		5 AMSLER	95C CBAR	$0.0 \bar{p} p \rightarrow \eta \eta \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1447 \pm 16 \pm 13$	163	6,7 DOBBS	15	$J/\psi \rightarrow \gamma \pi^+ \pi^-$
$1442 \pm 9 \pm 4$	261	6,7 DOBBS	15	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-$
$1486 \pm 10$		2 ANISOVICH	09 RVUE	$0.0 \bar{p} p, \pi N$
$1470 \pm 60$	568	8 KLEMP	08 E791	$D_S^+ \rightarrow \pi^- \pi^+ \pi^+$
$1470 \pm^{6+}_{-7} +^{72}_{-255}$		9 UEHARA	08A BELL	$10.6 e^+ e^- \rightarrow e^+ e^- \pi^0 \pi^0$
$1495 \pm 4$		AMSLER	06 CBAR	$0.9 \bar{p} p \rightarrow K^+ K^- \pi^0$
$1539 \pm 20$	9.9k	AUBERT	06O BABR	$B^+ \rightarrow K^+ K^- K^+$
$1473 \pm 5$	80k	10,11 UMAN	06 E835	$5.2 \bar{p} p \rightarrow \eta \eta \pi^0$
$1478 \pm 6$		VLADIMIRSK...	06 SPEC	$40 \pi^- p \rightarrow K_S^0 K_S^0 n$
$1493 \pm 7$		10 BINON	05 GAMS	$33 \pi^- p \rightarrow \eta \eta n$
$1524 \pm 14$	1400	12 GARMASH	05 BELL	$B^+ \rightarrow K^+ K^+ K^-$
$1489 \pm^{8+}_{-4}$		13 ANISOVICH	03 RVUE	
$1490 \pm 30$		10 ABELE	01 CBAR	$0.0 \bar{p} d \rightarrow \pi^- 4 \pi^0 p$
$1497 \pm 10$		10 BARBERIS	99 OMEG	$450 p p \rightarrow p_S p_f K^+ K^-$
$1502 \pm 10$		10 BARBERIS	99B OMEG	$450 p p \rightarrow p_S p_f \pi^+ \pi^-$
$1502 \pm 12 \pm 10$		14 BARBERIS	99D OMEG	$450 p p \rightarrow K^+ K^-, \pi^+ \pi^-$
$1530 \pm 45$		10 BELLAZZINI	99 GAM4	$450 p p \rightarrow p p \pi^0 \pi^0$
$1505 \pm 18$		10 FRENCH	99	$300 p p \rightarrow p_f (K^+ K^-) p_S$
$1447 \pm 27$		15 KAMINSKI	99 RVUE	$\pi \pi \rightarrow \pi \pi, K \bar{K}, \sigma \sigma$
$1580 \pm 80$		10 ALDE	98 GAM4	$100 \pi^- p \rightarrow \pi^0 \pi^0 n$
$1499 \pm 8$		2 ANISOVICH	98B RVUE	Compilation
$\sim 1520$		REYES	98 SPEC	$800 p p \rightarrow p_S p_f K_S^0 K_S^0$
$1510 \pm 20$		2 BARBERIS	97B OMEG	$450 p p \rightarrow p p 2(\pi^+ \pi^-)$
$\sim 1475$		FRABETTI	97D E687	$D_S^\pm \rightarrow \pi^\mp \pi^\pm \pi^\pm$
$\sim 1505$		ABELE	96 CBAR	$0.0 \bar{p} p \rightarrow 5 \pi^0$
$1500 \pm 8$		2 ABELE	96C RVUE	Compilation
$1460 \pm 20$	120	10 AMELIN	96B VES	$37 \pi^- A \rightarrow \eta \eta \pi^- A$
$1500 \pm 8$		BUGG	96 RVUE	
$1500 \pm 10$		16 AMSLER	95D CBAR	$0.0 \bar{p} p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta \eta, \pi^0 \pi^0 \eta$
$1445 \pm 5$		17 ANTINORI	95 OMEG	$300,450 p p \rightarrow p p 2(\pi^+ \pi^-)$
$1497 \pm 30$		10 ANTINORI	95 OMEG	$300,450 p p \rightarrow p p \pi^+ \pi^-$
$\sim 1505$		BUGG	95 MRK3	$J/\psi \rightarrow \gamma \pi^+ \pi^- \pi^+ \pi^-$
$1446 \pm 5$		10 ABATZIS	94 OMEG	$450 p p \rightarrow p p 2(\pi^+ \pi^-)$
$1545 \pm 25$		10 AMSLER	94E CBAR	$0.0 \bar{p} p \rightarrow \pi^0 \eta \eta'$
$1520 \pm 25$		2,18 ANISOVICH	94 CBAR	$0.0 \bar{p} p \rightarrow 3 \pi^0, \pi^0 \eta \eta$
$1505 \pm 20$		2,19 BUGG	94 RVUE	$\bar{p} p \rightarrow 3 \pi^0, \eta \eta \pi^0, \eta \pi^0 \pi^0$
$1560 \pm 25$		10 AMSLER	92 CBAR	$0.0 \bar{p} p \rightarrow \pi^0 \eta \eta$
$1550 \pm 45 \pm 30$		10 BELADIDZE	92C VES	$36 \pi^- \text{Be} \rightarrow \pi^- \eta' \eta \text{Be}$
$1449 \pm 4$		10 ARMSTRONG	89E OMEG	$300 p p \rightarrow p p 2(\pi^+ \pi^-)$
$1610 \pm 20$		10 ALDE	88 GAM4	$300 \pi^- N \rightarrow \pi^- N 2 \eta$
$\sim 1525$		ASTON	88D LASS	$11 K^- p \rightarrow K_S^0 K_S^0 \Lambda$
$1570 \pm 20$	600	10 ALDE	87 GAM4	$100 \pi^- p \rightarrow 4 \pi^0 n$
$1575 \pm 45$		20 ALDE	86D GAM4	$100 \pi^- p \rightarrow 2 \eta n$
$1568 \pm 33$		10 BINON	84C GAM2	$38 \pi^- p \rightarrow \eta \eta' n$
$1592 \pm 25$		10 BINON	83 GAM2	$38 \pi^- p \rightarrow 2 \eta n$
$1525 \pm 5$		10 GRAY	83 DBC	$0.0 \bar{p} N \rightarrow 3 \pi$

<sup>1</sup> From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.<sup>2</sup> T-matrix pole.<sup>3</sup> Average between  $\pi^+ \pi^- 2 \pi^0$  and  $2(\pi^+ \pi^-)$ .<sup>4</sup> T-matrix pole, supersedes ANISOVICH 94.<sup>5</sup> T-matrix pole, supersedes ANISOVICH 94 and AMSLER 92.<sup>6</sup> Using CLEO-c data but not authored by the CLEO Collaboration.<sup>7</sup> From a fit to a Breit-Wigner line shape with fixed  $\Gamma = 109$  MeV.<sup>8</sup> Reanalysis of AITALA 01A data. This state could also be  $f_0(1370)$ .<sup>9</sup> Breit-Wigner mass. May also be the  $f_0(1370)$ .<sup>10</sup> Breit-Wigner mass.<sup>11</sup> Statistical error only.

# Meson Particle Listings

## $f_0(1500)$

- <sup>12</sup> Breit-Wigner, solution 1, PWA ambiguous.
- <sup>13</sup> K-matrix pole from combined analysis of  $\pi^- p \rightarrow \pi^0 \pi^0 n$ ,  $\pi^- p \rightarrow K \bar{K} n$ ,  $\pi^+ \pi^- \rightarrow \pi^+ \pi^-$ ,  $\bar{p} p \rightarrow \pi^0 \pi^0 \pi^0$ ,  $\pi^0 \eta \eta$ ,  $\pi^0 \pi^0 \eta$ ,  $\pi^+ \pi^- \pi^0$ ,  $K^+ K^- \pi^0$ ,  $K_S^0 K_S^0 \pi^0$ ,  $K^+ K_S^0 \pi^-$  at rest,  $\bar{p} n \rightarrow \pi^- \pi^- \pi^+$ ,  $K_S^0 K^- \pi^0$ ,  $K_S^0 K_S^0 \pi^-$  at rest.
- <sup>14</sup> Supersedes BARBERIS 99 and BARBERIS 99b.
- <sup>15</sup> T-matrix pole on sheet  $--+$ .
- <sup>16</sup> T-matrix pole. Coupled-channel analysis of AMSLER 95b, AMSLER 95c, and AMSLER 94d.
- <sup>17</sup> Supersedes ABATZIS 94, ARMSTRONG 89e. Breit-Wigner mass.
- <sup>18</sup> From a simultaneous analysis of the annihilations  $\bar{p} p \rightarrow 3\pi^0, \pi^0 \eta \eta$ .
- <sup>19</sup> Reanalysis of ANISOVICH 94 data.
- <sup>20</sup> From central value and spread of two solutions. Breit-Wigner mass.



- 148<sup>+</sup><sub>-25</sub> 20 22,35 ANISOVICH 94 CBAR 0.0  $\bar{p} p \rightarrow 3\pi^0, \pi^0 \eta \eta$
- 150<sup>±</sup> 20 22,36 BUGG 94 RVUE  $\bar{p} p \rightarrow 3\pi^0, \eta \eta \pi^0, \eta \pi^0 \pi^0$
- 245<sup>±</sup> 50 27 AMSLER 92 CBAR 0.0  $\bar{p} p \rightarrow \pi^0 \eta \eta$
- 153<sup>±</sup> 67<sup>±</sup> 50 27 BELADIDZE 92c VES 36  $\pi^- \text{Be} \rightarrow \pi^- \eta' \eta \text{Be}$
- 78<sup>±</sup> 18 27 ARMSTRONG 89e OMEG 300  $p p \rightarrow p p 2(\pi^+ \pi^-)$
- 170<sup>±</sup> 40 27 ALDE 88 GAM4 300  $\pi^- N \rightarrow \pi^- N 2\eta$
- 150<sup>±</sup> 20 600 27 ALDE 87 GAM4 100  $\pi^- p \rightarrow 4\pi^0 n$
- 265<sup>±</sup> 65 37 ALDE 86d GAM4 100  $\pi^- p \rightarrow 2\eta n$
- 260<sup>±</sup> 60 27 BINON 84c GAM2 38  $\pi^- p \rightarrow \eta \eta' n$
- 210<sup>±</sup> 40 27 BINON 83 GAM2 38  $\pi^- p \rightarrow 2\eta n$
- 101<sup>±</sup> 13 27 GRAY 83 DBC 0.0  $\bar{p} N \rightarrow 3\pi$
- <sup>21</sup> From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.
- <sup>22</sup> T-matrix pole.
- <sup>23</sup> Average between  $\pi^+ \pi^- 2\pi^0$  and  $2(\pi^+ \pi^-)$ .
- <sup>24</sup> T-matrix pole, supersedes ANISOVICH 94.
- <sup>25</sup> T-matrix pole, supersedes ANISOVICH 94 and AMSLER 92.
- <sup>26</sup> Breit-Wigner width. May also be the  $f_0(1370)$ .
- <sup>27</sup> Breit-Wigner width.
- <sup>28</sup> Statistical error only.
- <sup>29</sup> Breit-Wigner, solution 1, PWA ambiguous.
- <sup>30</sup> K-matrix pole from combined analysis of  $\pi^- p \rightarrow \pi^0 \pi^0 n$ ,  $\pi^- p \rightarrow K \bar{K} n$ ,  $\pi^+ \pi^- \rightarrow \pi^+ \pi^-$ ,  $\bar{p} p \rightarrow \pi^0 \pi^0 \pi^0$ ,  $\pi^0 \eta \eta$ ,  $\pi^0 \pi^0 \eta$ ,  $\pi^+ \pi^- \pi^0$ ,  $K^+ K^- \pi^0$ ,  $K_S^0 K_S^0 \pi^0$ ,  $K^+ K_S^0 \pi^-$  at rest,  $\bar{p} n \rightarrow \pi^- \pi^- \pi^+$ ,  $K_S^0 K^- \pi^0$ ,  $K_S^0 K_S^0 \pi^-$  at rest.
- <sup>31</sup> Supersedes BARBERIS 99 and BARBERIS 99b.
- <sup>32</sup> T-matrix pole on sheet  $--+$ .
- <sup>33</sup> T-matrix pole. Coupled-channel analysis of AMSLER 95b, AMSLER 95c, and AMSLER 94d.
- <sup>34</sup> Supersedes ABATZIS 94, ARMSTRONG 89e. Breit-Wigner mass.
- <sup>35</sup> From a simultaneous analysis of the annihilations  $\bar{p} p \rightarrow 3\pi^0, \pi^0 \eta \eta$ .
- <sup>36</sup> Reanalysis of ANISOVICH 94 data.
- <sup>37</sup> From central value and spread of two solutions. Breit-Wigner mass.

## $f_0(1500)$ DECAY MODES

Mode		Fraction ( $\Gamma_i/\Gamma$ )	Scale factor
$\Gamma_1$	$\pi\pi$	(34.9 $\pm$ 2.3) %	1.2
$\Gamma_2$	$\pi^+\pi^-$	seen	
$\Gamma_3$	$2\pi^0$	seen	
$\Gamma_4$	$4\pi^0$	(49.5 $\pm$ 3.3) %	1.2
$\Gamma_5$	$2\pi^+2\pi^-$	seen	
$\Gamma_6$	$2(\pi\pi)$ s-wave	seen	
$\Gamma_7$	$\rho\rho$	seen	
$\Gamma_8$	$\pi(1300)\pi$	seen	
$\Gamma_9$	$a_1(1260)\pi$	seen	
$\Gamma_{10}$	$\eta\eta$	( 5.1 $\pm$ 0.9) %	1.4
$\Gamma_{11}$	$\eta\eta'(958)$	( 1.9 $\pm$ 0.8) %	1.7
$\Gamma_{12}$	$K\bar{K}$	( 8.6 $\pm$ 1.0) %	1.1
$\Gamma_{13}$	$\gamma\gamma$	not seen	
$\Gamma_{14}$			

## CONSTRAINED FIT INFORMATION

An overall fit to 6 branching ratios uses 10 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 11.4$  for 6 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_4$	-83			
$x_{11}$	11	-52		
$x_{12}$	-5	-31	29	
$x_{13}$	39	-67	33	6
$x_1$	$x_4$	$x_{11}$	$x_{12}$	

## $f_0(1500) \Gamma(\pi)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$		$\Gamma_1 \Gamma_{14}/\Gamma$	
VALUE (eV)	CL%	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
33 <sup>+</sup> 12 <sup>+</sup> 1809 <sup>+</sup> - 6 - 21		38 UEHARA 08A BELL	10.6 e <sup>+</sup> e <sup>-</sup> → e <sup>+</sup> e <sup>-</sup> π <sup>0</sup> π <sup>0</sup>
not seen		ACCIARRI 01H L3	γγ → K <sub>S</sub> <sup>0</sup> K <sub>S</sub> <sup>0</sup> , E <sub>cm</sub> <sup>ee</sup> = 91, 183-209 GeV
<460	95	BARATE 00E ALEP	γγ → π <sup>+</sup> π <sup>-</sup>
38 May also be the $f_0(1370)$ . Multiplied by us by 3 to obtain the $\pi\pi$ value.			



See key on page 601

## Meson Particle Listings

 $f_0(1500)$  $f_0(1500)$  BRANCHING RATIOS $\Gamma(\pi\pi)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN
0.454 ± 0.104	BUGG	96 RVUE

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
seen	BERTIN	98 OBLX	0.05–0.405 $\pi p \rightarrow \pi^+\pi^+\pi^-$
possibly seen	FRABETTI	97D E687	$D_S^\pm \rightarrow \pi^\mp\pi^\pm\pi^\pm$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(4\pi)/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.42 ± 0.18 OUR FIT</b>	Error includes scale factor of 1.2.		
<b>1.42 ± 0.18 OUR AVERAGE</b>	Error includes scale factor of 1.2.		
1.37 ± 0.16	BARBERIS	00D	450 $pp \rightarrow p_f 4\pi p_s$
2.1 ± 0.6	39 AMSLER	98 RVUE	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.1 ± 0.2	40 ANISOVICH	02D SPEC	Combined fit
3.4 ± 0.8	39 ABELE	96 CBAR	0.0 $\bar{p}p \rightarrow 5\pi^0$

 $\Gamma(2(\pi\pi)_{s\text{-wave}})/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
0.42 ± 0.26	41 ABELE	01 CBAR	0.0 $\bar{p}d \rightarrow \pi^- 4\pi^0 p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(2(\pi\pi)_{s\text{-wave}})/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
0.26 ± 0.07	ABELE	01B CBAR	0.0 $\bar{p}d \rightarrow 5\pi p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\rho\rho)/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
0.13 ± 0.08	ABELE	01B CBAR	0.0 $\bar{p}d \rightarrow 5\pi p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\rho\rho)/\Gamma(2(\pi\pi)_{s\text{-wave}})$ 

VALUE	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
3.3 ± 0.5	BARBERIS	00C 450 $pp \rightarrow p_f \pi^+ \pi^- 2\pi^0 p_s$
2.6 ± 0.4	BARBERIS	00C 450 $pp \rightarrow p_f 2(\pi^+ \pi^-) p_s$

 $\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
0.50 ± 0.25	ABELE	01B CBAR	0.0 $\bar{p}d \rightarrow 5\pi p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
0.12 ± 0.05	ABELE	01B CBAR	0.0 $\bar{p}d \rightarrow 5\pi p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\eta\eta)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
large	ALDE	88 GAM4	300 $\pi^- N \rightarrow \eta\eta\pi^- N$
large	BINON	83 GAM2	38 $\pi^- p \rightarrow 2\eta n$

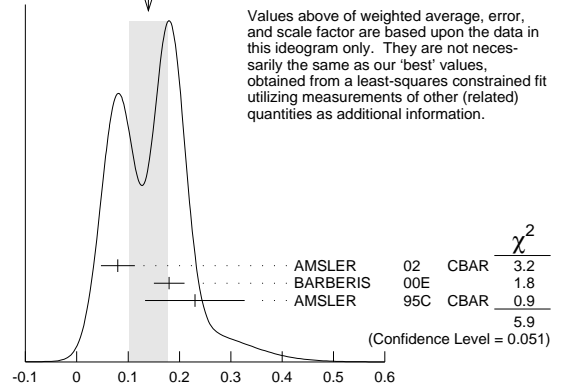
• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\eta\eta)/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.145 ± 0.027 OUR FIT</b>	Error includes scale factor of 1.5.		
<b>0.14 ± 0.04 OUR AVERAGE</b>	Error includes scale factor of 1.7. See the ideogram below.		
0.080 ± 0.033	AMSLER	02 CBAR	0.9 $\bar{p}p \rightarrow \pi^0\eta\eta, \pi^0\pi^0\pi^0$
0.18 ± 0.03	BARBERIS	00E	450 $pp \rightarrow p_f \eta\eta p_s$
0.230 ± 0.097	42 AMSLER	95C CBAR	0.0 $\bar{p}p \rightarrow \eta\eta\pi^0$

 $\Gamma_1/\Gamma$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.11 ± 0.03	40 ANISOVICH	02D SPEC	Combined fit	
0.078 ± 0.013	43 ABELE	96C RVUE	Compilation	
0.157 ± 0.060	44 AMSLER	95D CBAR	0.0 $\bar{p}p \rightarrow \pi^0\pi^0\pi^0, \pi^0\eta\eta, \pi^0\pi^0\eta$	

WEIGHTED AVERAGE  
0.14 ± 0.04 (Error scaled by 1.7) $\Gamma(4\pi^0)/\Gamma(\eta\eta)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
0.8 ± 0.3	ALDE	87 GAM4	100 $\pi^- p \rightarrow 4\pi^0 n$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\eta\eta'(958))/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.055 ± 0.024 OUR FIT</b>	Error includes scale factor of 1.8.		
<b>0.095 ± 0.026</b>	BARBERIS	00A	450 $pp \rightarrow p_f \eta\eta p_s$
0.005 ± 0.003	40 ANISOVICH	02D SPEC	Combined fit

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(\eta\eta'(958))/\Gamma(\eta\eta)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.38 ± 0.16 OUR FIT</b>	Error includes scale factor of 1.9.		
<b>0.29 ± 0.10</b>	45 AMSLER	95C CBAR	0.0 $\bar{p}p \rightarrow \eta\eta\pi^0$
0.05 ± 0.03	40 ANISOVICH	02D SPEC	Combined fit
0.84 ± 0.23	ABELE	96C RVUE	Compilation
2.7 ± 0.8	BINON	84C GAM2	38 $\pi^- p \rightarrow \eta\eta' n$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(K\bar{K})/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN
0.044 ± 0.021	BUGG	96 RVUE

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(K\bar{K})/\Gamma(\pi\pi)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.246 ± 0.026 OUR FIT</b>	Error includes scale factor of 1.4.		
<b>0.241 ± 0.028 OUR AVERAGE</b>			
0.25 ± 0.03	46 BARGIOTTI	03 OBLX	$\bar{p}p$
0.19 ± 0.07	47 ABELE	98 CBAR	0.0 $\bar{p}p \rightarrow K_L^0 K^\pm \pi^\mp$
0.16 ± 0.05	40 ANISOVICH	02D SPEC	Combined fit
0.33 ± 0.03 ± 0.07	BARBERIS	99D OMEG	450 $pp \rightarrow K^+ K^-, \pi^+ \pi^-$
0.20 ± 0.08	48 ABELE	96B CBAR	0.0 $\bar{p}p \rightarrow \pi^0 K_L^0 K_L^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\Gamma(K\bar{K})/\Gamma(\eta\eta)$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.69 ± 0.33 OUR FIT</b>	Error includes scale factor of 1.4.			
<b>1.85 ± 0.41</b>		BARBERIS	00E	450 $pp \rightarrow p_f \eta\eta p_s$
1.5 ± 0.6	40	ANISOVICH	02D SPEC	Combined fit
<0.4	90	49 PROKOSHIN	91 GAM4	300 $\pi^- p \rightarrow \pi^- \pi^- \rho\eta\eta$
<0.6	50	BINON	83 GAM2	38 $\pi^- p \rightarrow 2\eta n$

• • • We do not use the following data for averages, fits, limits, etc. • • •

39 Excluding  $\rho\rho$  contribution to  $4\pi$ .  
 40 From a combined K-matrix analysis of Crystal Barrel (0.  $\bar{p}p \rightarrow \pi^0\pi^0\pi^0, \pi^0\eta\eta, \pi^0\pi^0\eta$ ), GAMS ( $\pi p \rightarrow \pi^0\pi^0 n, \eta\eta n, \eta\eta' n$ ), and BNL ( $\pi p \rightarrow K\bar{K} n$ ) data.  
 41 From the combined data of ABELE 96 and ABELE 96c.  
 42 Using AMSLER 95B (3 $\pi^0$ ).

# Meson Particle Listings

## $f_0(1500)$ , $f_1(1510)$

- <sup>43</sup>  $2\pi$  width determined to be  $60 \pm 12$  MeV.  
<sup>44</sup> Coupled-channel analysis of AMSLER 95B, AMSLER 95C, and AMSLER 94D.  
<sup>45</sup> Using AMSLER 94E ( $\eta\eta'\pi^0$ ).  
<sup>46</sup> Coupled channel analysis of  $\pi^+\pi^-\pi^0$ ,  $K^+K^-\pi^0$ , and  $K^\pm K_S^0 \pi^\mp$ .  
<sup>47</sup> Using  $\pi^0\pi^0$  from AMSLER 95B.  
<sup>48</sup> Using AMSLER 95B ( $3\pi^0$ ), AMSLER 94C ( $2\pi^0\eta$ ) and SU(3).  
<sup>49</sup> Combining results of GAM4 with those of WA76 on  $K\bar{K}$  central production.  
<sup>50</sup> Using ETKIN 82B and COHEN 80.

### $f_0(1500)$ REFERENCES

DOBBS	15	PR D91 052006	S. Dobbs <i>et al.</i>	(NWES)
ABLIKIM	13N	PR D87 092009	Ablikim M. <i>et al.</i>	(BES III Collab.)
ANISOVICH	09	UHP A24 2481	V.V. Anisovich, A.V. Sarantsev	
KLEMPT	08	EPJ C55 39	E. Klempf, M. Matveev, A.V. Sarantsev	(BONN+)
UEHARA	08A	PR D78 052004	S. Uehara <i>et al.</i>	(BELLE Collab.)
ABLIKIM	06V	PL B642 441	M. Ablikim <i>et al.</i>	(BES Collab.)
AMSLER	06	PL B639 165	C. Amsler <i>et al.</i>	(CBAR Collab.)
AUBERT	06O	PR D74 032003	B. Aubert <i>et al.</i>	(BABAR Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
UMAN	06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
VLADIMIRSK...	06	PAN 69 493	V.V. Vladimirovsky <i>et al.</i>	(ITEP, Moscow)
Translated from YAF 69 515.				
BINON	05	PAN 68 960	F. Binon <i>et al.</i>	
Translated from YAF 68 998.				
GARMASH	05	PR D71 092003	A. Garmash <i>et al.</i>	(BELLE Collab.)
ANISOVICH	03	EPJ A16 229	V.V. Anisovich <i>et al.</i>	
BARGIOTTI	03	EPJ C26 371	M. Bargiotti <i>et al.</i>	(OBELIX Collab.)
AMSLER	02	EPJ C23 29	C. Amsler <i>et al.</i>	
ANISOVICH	02D	PAN 65 1545	V.V. Anisovich <i>et al.</i>	
Translated from YAF 65 1583.				
ABELE	01	EPJ C19 667	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	01B	EPJ C21 261	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACCIARRI	01H	PL B501 173	M. Acciarri <i>et al.</i>	(L3 Collab.)
AITALA	01A	PRL 86 765	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BARATE	00E	PL B472 189	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARBERIS	00A	PL B471 429	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00C	PL B471 440	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00D	PL B474 423	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00E	PL B479 59	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	99	PL B453 305	D. Barberis <i>et al.</i>	(Omega Expt.)
BARBERIS	99B	PL B453 316	D. Barberis <i>et al.</i>	(Omega Expt.)
BARBERIS	99D	PL B462 462	D. Barberis <i>et al.</i>	(Omega Expt.)
BELLAZZINI	99	PL B467 296	R. Bellazzini <i>et al.</i>	(Omega Expt.)
FRENCH	99	PL B460 213	B. French <i>et al.</i>	(WA76 Collab.)
KAMINSKI	99	EPJ C9 141	R. Kaminski, L. Lesniak, B. Loiseau	(CRAC, PARIN)
ABELE	98	PR D57 3860	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ALDE	98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
Also		PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
Translated from YAF 62 446.				
AMSLER	98	RMP 70 1293	C. Amsler	
ANISOVICH	98B	SPU 41 419	V.V. Anisovich <i>et al.</i>	
Translated from UFN 168 481.				
BERTIN	98	PR D57 55	A. Bertin <i>et al.</i>	(OBELIX Collab.)
REYES	98	PRL 81 4079	M.A. Reyes <i>et al.</i>	
BARBERIS	97B	PL B413 217	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
FRABETTI	97D	PL B407 79	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ABELE	96	PL B380 453	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	96B	PL B385 425	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	96C	NP A609 562	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
AMELIN	96B	PAN 59 976	D.V. Amelin <i>et al.</i>	(SERP, TBIL)
Translated from YAF 59 1021.				
BUGG	96	NP B471 59	D.V. Bugg, A.V. Sarantsev, B.S. Zou	(LOQM, PNPI)
AMSLER	95B	PL B342 433	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	95C	PL B353 571	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	95D	PL B355 425	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
ANTINORI	95	PL B353 589	F. Antinori <i>et al.</i>	(ATHU, BARI, BIRM+)
BUGG	95	PL B353 378	D.V. Bugg <i>et al.</i>	(LOQM, PNPI, WASH)
ABATZIS	94	PL B324 509	S. Abatzis <i>et al.</i>	(ATHU, BARI, BIRM+)
AMSLER	94C	PL B327 425	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	94D	PL B333 277	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
AMSLER	94E	PL B340 259	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
ANISOVICH	94	PL B323 233	V.V. Anisovich <i>et al.</i>	(Crystal Barrel Collab.)
BUGG	94	PR D50 4412	D.V. Bugg <i>et al.</i>	(LOQM)
AMSLER	92	PL B291 347	C. Amsler <i>et al.</i>	(Crystal Barrel Collab.)
BELADIDZE	92C	SJNP 55 1535	G.M. Beladidze, S.I. Bityukov, G.V. Borisov	(SERP+)
Translated from YAF 55 2748.				
PROKOSHKIN	91	SPD 36 155	V.D. Prokoshkin	(GAM2, GAM4 Collab.)
Translated from DANS 316 900.				
ARMSTRONG	89E	PL B228 536	T.A. Armstrong, M. Benayoun	(ATHU, BARI, BIRM+)
ALDE	88	PL B201 160	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP+)
ASTON	88D	NP B301 525	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ALDE	87	PL B198 286	D.M. Alde <i>et al.</i>	(LANL, BRUX, SERP, LAPP)
ALDE	86D	NP B269 485	D.M. Alde <i>et al.</i>	(BELG, LAPP, SERP, CERN+)
BINON	84C	NC 80A 363	F.G. Binon <i>et al.</i>	(BELG, LAPP, SERP+)
BINON	83	NC 78A 313	F.G. Binon <i>et al.</i>	(BELG, LAPP, SERP+)
Also		SJNP 38 561	F.G. Binon <i>et al.</i>	(BELG, LAPP, SERP+)
Translated from YAF 38 934.				
GRAY	83	PR D27 307	L. Gray <i>et al.</i>	(SYRA)
ETKIN	82B	PR D25 1786	A. Etkin <i>et al.</i>	(BNL, CUNY, TUFTS, VAND)
COHEN	80	PR D22 2595	D. Cohen <i>et al.</i>	(ANL)

### $f_1(1510)$

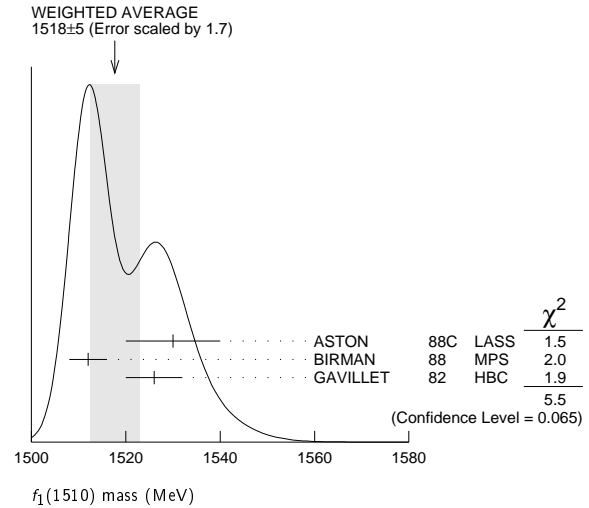
$$I^G(J^{PC}) = 0^+(1^{++})$$

OMITTED FROM SUMMARY TABLE  
See the minireview under  $\eta(1405)$ .

### $f_1(1510)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1518± 5 OUR AVERAGE</b>		Error includes scale factor of 1.7. See the ideogram below.		
1530±10		ASTON	88C LASS	$11 K^-p \rightarrow K_S^0 K^\pm \pi^\mp \Lambda$
1512± 4	600	<sup>1</sup> BIRMAN	88C MPS	$8 \pi^-p \rightarrow K^+ \bar{K}^0 \pi^- n$
1526± 6	271	GAVILLET	82 HBC	$4.2 K^-p \rightarrow \Lambda K K \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 1525		<sup>2</sup> BAUER	93B	$\gamma \gamma^* \rightarrow \pi^+ \pi^- \pi^0 \pi^0$

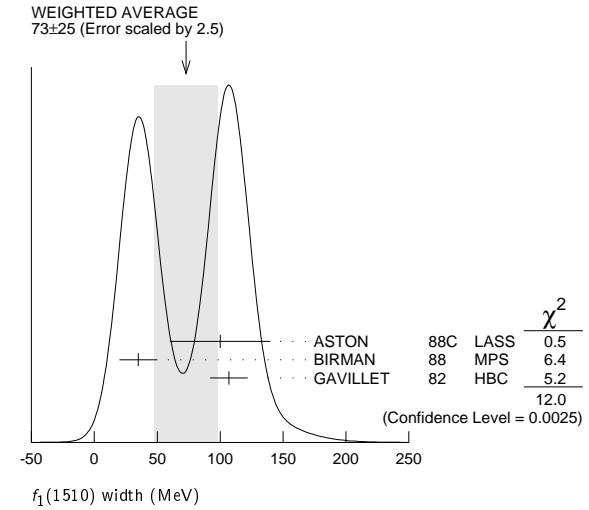
- <sup>1</sup> From partial wave analysis of  $K^+ \bar{K}^0 \pi^-$  state.  
<sup>2</sup> Not seen by AIHARA 88C in the  $K_S^0 K^\pm \pi^\mp$  final state.



### $f_1(1510)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>73±25 OUR AVERAGE</b>	Error	includes scale factor of 2.5. See the ideogram below.		
100±40		ASTON	88C LASS	$11 K^-p \rightarrow K_S^0 K^\pm \pi^\mp \Lambda$
35±15	600	<sup>3</sup> BIRMAN	88 MPS	$8 \pi^-p \rightarrow K^+ \bar{K}^0 \pi^- n$
107±15	271	GAVILLET	82 HBC	$4.2 K^-p \rightarrow \Lambda K K \pi$

<sup>3</sup> From partial wave analysis of  $K^+ \bar{K}^0 \pi^-$  state.



### $f_1(1510)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K \bar{K}^*(892) + c.c.$	seen
$\Gamma_2$ $\pi^+ \pi^- \eta'$	seen

### $f_1(1510)$ BRANCHING RATIOS

$\Gamma(\pi^+ \pi^- \eta')/\Gamma_{total}$				$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	230	ABL IKIM	11c BES3	$J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$

### $f_1(1510)$ REFERENCES

ABL IKIM	11C	PRL 106 072002	M. Ablikim <i>et al.</i>	(BES III Collab.)
BAUER	93B	PR D48 3976	D.A. Bauer <i>et al.</i>	(SLAC)
AIHARA	88C	PR D38 1	H. Aihara <i>et al.</i>	(TPC-2γ Collab.)
ASTON	88C	PL B201 573	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)JP
BIRMAN	88	PRL 61 1557	A. Birman <i>et al.</i>	(BNL, FSU, IND, MASD)JP
GAVILLET	82	ZPHY C16 119	P. Gavillet <i>et al.</i>	(CERN, CDEF, PADO+)

See key on page 601

## Meson Particle Listings

 $f'_2(1525)$ 

$f'_2(1525)$	$J^G(J^{PC}) = 0^+(2^{++})$
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 $f'_2(1525)$  MASS

VALUE (MeV)	DOCUMENT ID
<b>1525 ± 5 OUR ESTIMATE</b>	This is only an educated guess; the error given is larger than the error on the average of the published values.

## PRODUCED BY PION BEAM

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1521 ± 13		TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
1547 $^{+10}_{-2}$		<sup>1</sup> LONGACRE 86	MPS	22 $\pi^- p \rightarrow K_S^0 K_S^0 n$
1496 $^{+9}_{-8}$		<sup>2</sup> CHABAUD 81	ASPK	6 $\pi^- p \rightarrow K^+ K^- n$
1497 $^{+8}_{-9}$		CHABAUD 81	ASPK	18.4 $\pi^- p \rightarrow K^+ K^- n$
1492 ± 29		GORLICH 80	ASPK	17 $\pi^- p$ polarized $\rightarrow K^+ K^- n$
1502 ± 25		<sup>3</sup> CORDEN 79	OMEG	12–15 $\pi^- p \rightarrow \pi^+ \pi^- n$
1480	14	CRENNELL 66	HBC	6.0 $\pi^- p \rightarrow K_S^0 K_S^0 n$

PRODUCED BY  $K^\pm$  BEAM

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1523.3± 1.1 OUR AVERAGE</b>	Includes data from the datablock that follows this one.			
Error includes scale factor of 1.1.				
1526.8± 4.3		ASTON 88D	LASS	11 $K^-p \rightarrow K_S^0 K_S^0 \Lambda$
1504 ± 12		BOLONKIN 86	SPEC	40 $K^-p \rightarrow K_S^0 K_S^0 Y$
1529 ± 3		ARMSTRONG 83B	OMEG	18.5 $K^-p \rightarrow K^- K^+ \Lambda$
1521 ± 6	650	AGUILAR... 81B	HBC	4.2 $K^-p \rightarrow \Lambda K^+ K^-$
1521 ± 3	572	ALHARRAN 81	HBC	8.25 $K^-p \rightarrow \Lambda K \bar{K}$
1522 ± 6	123	BARREIRO 77	HBC	4.15 $K^-p \rightarrow \Lambda K_S^0 K_S^0$
1528 ± 7	166	EVANGELIS... 77	OMEG	10 $K^-p \rightarrow K^+ K^- (\Lambda, \Sigma)$
1527 ± 3	120	BRANDENB... 76C	ASPK	13 $K^-p \rightarrow K^+ K^- (\Lambda, \Sigma)$
1519 ± 7	100	AGUILAR... 72B	HBC	3.9,4.6 $K^-p \rightarrow K \bar{K} (\Lambda, \Sigma)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1514 ± 8	61	BINON 07	GAMS	32.5 $K^-p \rightarrow \eta \eta (\Lambda / \Sigma^0)$
1513 ± 10	<sup>4</sup>	BARKOV 99	SPEC	40 $K^-p \rightarrow K_S^0 K_S^0 y$

PRODUCED IN  $e^+ e^-$  ANNIHILATION AND PARTICLE DECAYS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.				

**1521.9 ± 1.8 OUR AVERAGE** Error includes scale factor of 1.1.

1522.2 ± 2.8 $^{+5.3}_{-2.0}$		AAIJ	13AN LHCB	$\bar{B}_s^0 \rightarrow J/\psi K^+ K^-$
1513 ± 5 $^{+4}_{-10}$	5.5k	<sup>5</sup> ABLIKIM	13N BES3	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma \eta \eta$
1525.3 $^{+1.2}_{-1.4}$ $^{+3.7}_{-2.1}$		UEHARA	13 BELL	$\gamma \gamma \rightarrow K_S^0 K_S^0$
1521 ± 5		ABLIKIM	05 BES2	$J/\psi \rightarrow \phi K^+ K^-$
1518 ± 1 ± 3		ABE	04 BELL	10.6 $e^+ e^- \rightarrow K^+ K^-$ $e^+ e^- K^+ K^-$
1519 ± 2 $^{+15}_{-5}$		BAI	03G BES	$J/\psi \rightarrow \gamma K \bar{K}$
1523 ± 6	331	<sup>6</sup> ACCIARRI	01H L3	91, 183–209 $e^+ e^- \rightarrow$ $e^+ e^- K_S^0 K_S^0$
1535 ± 5 ± 4		ABREU	96C DLPH	$Z^0 \rightarrow K^+ K^- X$
1516 ± 5 $^{+9}_{-15}$		BAI	96C BES	$J/\psi \rightarrow \gamma K^+ K^-$
1531.6 ± 10.0		AUGUSTIN	88 DM2	$J/\psi \rightarrow \gamma K^+ K^-$
1515 ± 5		<sup>7</sup> FALVARD	88 DM2	$J/\psi \rightarrow \phi K^+ K^-$
1525 ± 10 ± 10		BALTRUSAIT... 87	MRK3	$J/\psi \rightarrow \gamma K^+ K^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1532 ± 3 ± 6	644	<sup>8,9</sup> DOBBS	15	$J/\psi \rightarrow \gamma K^+ K^-$
1557 ± 9 ± 3	113	<sup>8,9</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K^+ K^-$
1523 ± 5	870	<sup>10</sup> SCHEGELSKY	06A RVUE	$\gamma \gamma \rightarrow K_S^0 K_S^0$
1496 ± 2		<sup>11</sup> FALVARD	88 DM2	$J/\psi \rightarrow \phi K^+ K^-$

PRODUCED IN  $\bar{p} p$  ANNIHILATION

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1530 ± 12	<sup>12</sup> ANISOVICH 09	RVUE	0.0 $\bar{p} p, \pi N$
1513 ± 4	AMSLER 06	CBAR	0.9 $\bar{p} p \rightarrow K^+ K^- \pi^0$
1508 ± 9	<sup>13</sup> AMSLER 02	CBAR	0.9 $\bar{p} p \rightarrow \pi^0 \eta \eta, \pi^0 \pi^0 \pi^0$

## CENTRAL PRODUCTION

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1515 ± 15</b>	BARBERIS 99	OMEG	450 $p p \rightarrow p_S p_f K^+ K^-$

PRODUCED IN  $ep$  COLLISIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1512 ± 3 <math>^{+1.4}_{-0.5}</math></b>		<sup>14</sup> CHEKANOV 08	ZEUS	$ep \rightarrow K_S^0 K_S^0 X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1537 $^{+9}_{-8}$	84	<sup>15</sup> CHEKANOV 04	ZEUS	$ep \rightarrow K_S^0 K_S^0 X$
<sup>1</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles.				
<sup>2</sup> CHABAUD 81 is a reanalysis of PAWLICKI 77 data.				
<sup>3</sup> From an amplitude analysis where the $f'_2(1525)$ width and elasticity are in complete disagreement with the values obtained from $K \bar{K}$ channel, making the solution dubious.				
<sup>4</sup> Systematic errors not estimated.				
<sup>5</sup> From partial wave analysis including all possible combinations of $0^{++}$ , $2^{++}$ , and $4^{++}$ resonances.				
<sup>6</sup> Supersedes ACCIARRI 95J.				
<sup>7</sup> From an analysis ignoring interference with $f_0(1710)$ .				
<sup>8</sup> Using CLEO-c data but not authored by the CLEO Collaboration.				
<sup>9</sup> From a fit to a Breit-Wigner line shape with fixed $\Gamma = 73$ MeV.				
<sup>10</sup> From analysis of L3 data at 91 and 183–209 GeV.				
<sup>11</sup> From an analysis including interference with $f_0(1710)$ .				
<sup>12</sup> 4-poles, 5-channel K matrix fit.				
<sup>13</sup> T-matrix pole.				
<sup>14</sup> In the SU(3) based model with a specific interference pattern of the $f_2(1270)$ , $a_2^0(1320)$ , and $f'_2(1525)$ mesons incoherently added to the $f_0(1710)$ and non-resonant background.				
<sup>15</sup> Systematic errors not estimated.				

 $f'_2(1525)$  WIDTH

VALUE (MeV)	DOCUMENT ID	COMMENT
<b>73 <math>^{+6}_{-5}</math> OUR FIT</b>		
<b>76 ± 10</b>	PDG 90	For fitting

## PRODUCED BY PION BEAM

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
102 ± 42	TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
108 $^{+5}_{-2}$	<sup>16</sup> LONGACRE 86	MPS	22 $\pi^- p \rightarrow K_S^0 K_S^0 n$
69 $^{+22}_{-16}$	<sup>17</sup> CHABAUD 81	ASPK	6 $\pi^- p \rightarrow K^+ K^- n$
137 $^{+23}_{-21}$	CHABAUD 81	ASPK	18.4 $\pi^- p \rightarrow K^+ K^- n$
150 $^{+83}_{-50}$	GORLICH 80	ASPK	17 $\pi^- p$ polarized $\rightarrow K^+ K^- n$
165 ± 42	<sup>18</sup> CORDEN 79	OMEG	12–15 $\pi^- p \rightarrow \pi^+ \pi^- n$
92 $^{+39}_{-22}$	<sup>19</sup> POLYCHRO... 79	STRC	7 $\pi^- p \rightarrow n K_S^0 K_S^0$

PRODUCED BY  $K^\pm$  BEAM

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>81.4<sup>+2.2</sup><sub>-1.9</sub> OUR AVERAGE</b>		Includes data from the datablock that follows this one.		
90 ±12		ASTON 88D	LASS	11 K <sup>-</sup> p → K <sub>S</sub> <sup>0</sup> K <sub>S</sub> <sup>0</sup> Λ
73 ±18		BOLONKIN 86	SPEC	40 K <sup>-</sup> p → K <sub>S</sub> <sup>0</sup> K <sub>S</sub> <sup>0</sup> Y
83 ±15		ARMSTRONG 83B	OMEG	18.5 K <sup>-</sup> p → K <sup>-</sup> K <sup>+</sup> Λ
85 ±16	650	AGUILAR... 81B	HBC	4.2 K <sup>-</sup> p → ΛK <sup>+</sup> K <sup>-</sup>
80 <sup>+14</sup> <sub>-11</sub>	572	ALHARRAN 81	HBC	8.25 K <sup>-</sup> p → ΛK $\overline{K}$
72 ±25	166	EVANGELIS... 77	OMEG	10 K <sup>-</sup> p → K <sup>+</sup> K <sup>-</sup> (Λ,Σ)
69 ±22	100	AGUILAR... 72B	HBC	3.9,4.6 K <sup>-</sup> p → K $\overline{K}$ (Λ,Σ)
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
92 <sup>+25</sup> <sub>-16</sub>	61	BINON 07	GAMS	32.5 K <sup>-</sup> p → ηη(Λ/Σ <sup>0</sup> )
75 ±20		<sup>20</sup> BARKOV 99	SPEC	40 K <sup>-</sup> p → K <sub>S</sub> <sup>0</sup> K <sub>S</sub> <sup>0</sup> γ
62 <sup>+19</sup> <sub>-14</sub>	123	BARREIRO 77	HBC	4.15 K <sup>-</sup> p → ΛK <sub>S</sub> <sup>0</sup> K <sub>S</sub> <sup>0</sup>
61 ± 8	120	BRANDENB... 76C	ASPK	13 K <sup>-</sup> p → K <sup>+</sup> K <sup>-</sup> (Λ,Σ)

PRODUCED IN  $e^+ e^-$  ANNIHILATION AND PARTICLE DECAYS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.				

**81.4  $^{+2.4}_{-2.0}$  OUR AVERAGE**

84 ± 6 $^{+10}_{-5}$		AAIJ	13AN LHCB	$\bar{B}_s^0 \rightarrow J/\psi K^+ K^-$
75 $^{+12}_{-10}$ $^{+16}_{-8}$	5.5k	<sup>21</sup> ABLIKIM	13N BES3	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma \eta \eta$
82.9 $^{+2.1}_{-2.2}$ $^{+3.3}_{-2.0}$		UEHARA	13 BELL	$\gamma \gamma \rightarrow K_S^0 K_S^0$
77 ± 15		ABLIKIM	05 BES2	$J/\psi \rightarrow \phi K^+ K^-$
82 ± 2 ± 3		ABE	04 BELL	10.6 $e^+ e^- \rightarrow$ $e^+ e^- K^+ K^-$
75 ± 4 $^{+15}_{-5}$		BAI	03G BES	$J/\psi \rightarrow \gamma K \bar{K}$

# Meson Particle Listings

## $f_2'(1525)$

100 ±15	331	22	ACCIARRI	01H	L3	91, 183–209 $e^+e^- \rightarrow e^+e^-K_S^0K_S^0$
60 ±20 ±19			ABREU	96c	DLPH	$Z^0 \rightarrow K^+K^- + X$
60 ±23 ±20 <sup>+13</sup> <sub>-20</sub>			BAI	96c	BES	$J/\psi \rightarrow \gamma K^+K^-$
103 ±30			AUGUSTIN	88	DM2	$J/\psi \rightarrow \gamma K^+K^-$
62 ±10		23	FALVARD	88	DM2	$J/\psi \rightarrow \phi K^+K^-$
85 ±35			BALTRUSAITIS	87	MRK3	$J/\psi \rightarrow \gamma K^+K^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
104 ±10	870	24	SCHEGELSKY	06A	RVUE	$\gamma\gamma \rightarrow K_S^0K_S^0$
100 ± 3		25	FALVARD	88	DM2	$J/\psi \rightarrow \phi K^+K^-$

### PRODUCED IN $\bar{p}p$ ANNIHILATION

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>79 ± 8</b>	26	AMSLER	02 CBAR 0.9 $\bar{p}p \rightarrow \pi^0\eta\eta, \pi^0\pi^0\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
128 ±20	27	ANISOVICH	09 RVUE 0.0 $\bar{p}p, \pi N$
76 ± 6		AMSLER	06 CBAR 0.9 $\bar{p}p \rightarrow K^+K^-\pi^0$

### CENTRAL PRODUCTION

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>70 ±25</b>		BARBERIS	99 OMEG 450 $pp \rightarrow p_S p_f K^+K^-$

### PRODUCED IN $e\bar{p}$ COLLISIONS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>83 ± 9<sup>+5</sup><sub>-4</sub></b>		28	CHEKANOV	08 ZEUS $e\bar{p} \rightarrow K_S^0K_S^0X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
50 <sup>+34</sup> <sub>-22</sub>		84	29	CHEKANOV 04 ZEUS $e\bar{p} \rightarrow K_S^0K_S^0X$

- 16 From a partial-wave analysis of data using a K-matrix formalism with 5 poles.  
17 CHABAUD 81 is a reanalysis of PAWLICKI 77 data.  
18 From an amplitude analysis where the  $f_2'(1525)$  width and elasticity are in complete disagreement with the values obtained from  $K\bar{K}$  channel, making the solution dubious.  
19 From a fit to the  $D$  with  $f_2(1270)$ - $f_2'(1525)$  interference. Mass fixed at 1516 MeV.  
20 Systematic errors not estimated.  
21 From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.  
22 Supersedes ACCIARRI 95j.  
23 From an analysis ignoring interference with  $f_0(1710)$ .  
24 From analysis of L3 data at 91 and 183–209 GeV.  
25 From an analysis including interference with  $f_0(1710)$ .  
26 T-matrix pole.  
27 4-poles, 5-channel K matrix fit.  
28 In the SU(3) based model with a specific interference pattern of the  $f_2(1270)$ ,  $a_2^0(1320)$ , and  $f_2'(1525)$  mesons incoherently added to the  $f_0(1710)$  and non-resonant background.  
29 Systematic errors not estimated.

### $f_2'(1525)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\bar{K}$	(88.7 ±2.2 ) %
$\Gamma_2$ $\eta\eta$	(10.4 ±2.2 ) %
$\Gamma_3$ $\pi\pi$	( 8.2 ±1.5 ) × 10 <sup>−3</sup>
$\Gamma_4$ $K\bar{K}^*(892) + \text{c.c.}$	
$\Gamma_5$ $\pi K\bar{K}$	
$\Gamma_6$ $\pi\pi\eta$	
$\Gamma_7$ $\pi^+\pi^+\pi^-\pi^-$	
$\Gamma_8$ $\gamma\gamma$	( 1.10 ±0.14 ) × 10 <sup>−6</sup>

### CONSTRAINED FIT INFORMATION

An overall fit to the total width, 2 partial widths, a combination of partial widths obtained from integrated cross sections, and 3 branching ratios uses 17 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 14.3$  for 13 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−100			
$x_3$	−6	−1		
$x_8$	−6	6	1	
$\Gamma$	−23	23	−1	−56
	$x_1$	$x_2$	$x_3$	$x_8$
Mode	Rate (MeV)			
$\Gamma_1$ $K\bar{K}$	65	<sup>+5</sup> <sub>−4</sub>		
$\Gamma_2$ $\eta\eta$	7.6	±1.8		

$\Gamma_3$ $\pi\pi$	0.60 ±0.12
$\Gamma_8$ $\gamma\gamma$	( 8.1 ±0.9 ) × 10 <sup>−5</sup>

### $f_2'(1525)$ PARTIAL WIDTHS

$\Gamma(K\bar{K})$				$\Gamma_1$
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	

**65 ±5<sup>+5</sup><sub>-4</sub> OUR FIT**

**63 ±6<sup>+6</sup><sub>-5</sub>**

30 LONGACRE 86 MPS 22  $\pi^-p \rightarrow K_S^0K_S^0n$

### $\Gamma(\eta\eta)$

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>7.6 ±1.8 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

5.0 ±0.8 870 31 SCHEGELSKY 06A RVUE  $\gamma\gamma \rightarrow K_S^0K_S^0$

24<sup>+3</sup><sub>−1</sub> 30 LONGACRE 86 MPS 22  $\pi^-p \rightarrow K_S^0K_S^0n$

### $\Gamma(\pi\pi)$

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.60 ±0.12 OUR FIT</b>				

1.4<sup>+1.0</sup><sub>−0.5</sub> 30 LONGACRE 86 MPS 22  $\pi^-p \rightarrow K_S^0K_S^0n$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.2<sup>+1.0</sup><sub>−0.2</sub> 870 31 SCHEGELSKY 06A RVUE  $\gamma\gamma \rightarrow K_S^0K_S^0$

### $\Gamma(\gamma\gamma)$

VALUE (keV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.081 ±0.009 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.13 ±0.03 870 31 SCHEGELSKY 06A RVUE  $\gamma\gamma \rightarrow K_S^0K_S^0$

30 From a partial-wave analysis of data using a K-matrix formalism with 5 poles.

31 From analysis of L3 data at 91 and 183–209 GeV, using  $\Gamma(f_2'(1525) \rightarrow K\bar{K}) = 68$  MeV and SU(3) relations.

### $f_2'(1525)$ $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_1\Gamma_8/\Gamma$
VALUE (keV)	EVTs	DOCUMENT ID	TECN	COMMENT	

**0.072 ±0.007 OUR FIT**

**0.072 ±0.007 OUR AVERAGE**

0.048<sup>+0.067</sup><sub>−0.008</sub><sup>+0.108</sup><sub>−0.012</sub> UEHARA 13 BELL  $\gamma\gamma \rightarrow K_S^0K_S^0$

0.0564 ±0.0048 ±0.0116 ABE 04 BELL 10.6  $e^+e^- \rightarrow e^+e^-K^+K^-$

0.076 ±0.006 ±0.011 331 32 ACCIARRI 01H L3  $e^+e^- \rightarrow e^+e^-K^+K^-K_S^0K_S^0$

0.067 ±0.008 ±0.015 33 ALBRECHT 90G ARG  $e^+e^- \rightarrow e^+e^-K^+K^-$

0.11<sup>+0.03</sup><sub>−0.02</sub> ±0.02 BEHREND 89c CELL  $e^+e^- \rightarrow e^+e^-K_S^0K_S^0$

0.10<sup>+0.04</sup><sub>−0.03</sub><sup>+0.03</sup><sub>−0.02</sub> BERGER 88 PLUT  $e^+e^- \rightarrow e^+e^-K_S^0K_S^0$

0.12 ±0.07 ±0.04 33 AIHARA 86B TPC  $e^+e^- \rightarrow e^+e^-K^+K^-$

0.11 ±0.02 ±0.04 33 ALTHOFF 83 TASS  $e^+e^- \rightarrow e^+e^-K\bar{K}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.0314 ±0.0050 ±0.0077 34 ALBRECHT 90G ARG  $e^+e^- \rightarrow e^+e^-K^+K^-$

32 Supersedes ACCIARRI 95j. From analysis of L3 data at 91 and 183–209 GeV,

33 Using an incoherent background.

34 Using a coherent background.

### $f_2'(1525)$ BRANCHING RATIOS

$\Gamma(\eta\eta)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	

• • • We do not use the following data for averages, fits, limits, etc. • • •

seen UEHARA 10A BELL 10.6  $e^+e^- \rightarrow e^+e^-\eta\eta$

0.10 ±0.03 35 PROKOSHKIN 91 GAM4 300  $\pi^-p \rightarrow \pi^-p\eta\eta$

35 Combining results of GAM4 with those of WA76 on  $K\bar{K}$  central production and results of CBAL, MRK3 and DM2 on  $J/\psi \rightarrow \gamma\eta\eta$ .

### $\Gamma(\eta\eta)/\Gamma(K\bar{K})$

VALUE	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.118 ±0.028 OUR FIT</b>					
<b>0.115 ±0.028 OUR AVERAGE</b>					

0.119 ±0.015 ±0.036 61 36 BINON 07 GAMS 32.5  $K^-p \rightarrow \eta\eta(\Lambda/\Sigma^0)$

0.11 ±0.04 37 PROKOSHKIN 91 GAM4 300  $\pi^-p \rightarrow \pi^-p\eta\eta$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.14 90 BARBERIS 00E 450  $p\bar{p} \rightarrow p_f\eta\eta p_S$

< 0.50 BARNES 67 HBC 4.6, 5.0  $K^-p$

36 Using the compilation of the cross sections for  $f_2'(1525)$  production in  $K^-p$  collisions from ASTON 88d.

37 Combining results of GAM4 with those of WA76 on  $K\bar{K}$  central production and results of CBAL, MRK3 and DM2 on  $J/\psi \rightarrow \gamma\eta\eta$ .

See key on page 601

## Meson Particle Listings

 $f'_2(1525), f_2(1565)$ 

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
<b>0.0082±0.0016 OUR FIT</b>					
<b>0.0075±0.0016 OUR AVERAGE</b>					
0.007 ±0.002		COSTA...	80	OMEG 10 $\pi^- p \rightarrow K^+ K^- n$	
0.027 $\pm^{+0.071}_{-0.013}$		<sup>38</sup> GORLICH	80	ASPK 17,18 $\pi^- p$	
0.0075±0.0025		<sup>38,39</sup> MARTIN	79	RVUE	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.06	95	AGUILAR...	81B	HBC 4.2 $K^- p \rightarrow \Lambda K^+ K^-$	
0.19 ±0.03		CORDEN	79	OMEG 12-15 $\pi^- p \rightarrow \pi^+ \pi^- n$	
<0.045	95	BARREIRO	77	HBC 4.15 $K^- p \rightarrow \Lambda K_S^0 K_S^0$	
0.012 ±0.004		<sup>38</sup> PAWLICKI	77	SPEC 6 $\pi N \rightarrow K^+ K^- N$	
<0.063	90	BRANDENB...	76C	ASPK 13 $K^- p \rightarrow K^+ K^- (\Lambda, \Sigma)$	
<0.0086		<sup>38</sup> BEUSCH	75B	OSPK 8.9 $\pi^- p \rightarrow K^0 \bar{K}^0 n$	

<sup>38</sup> Assuming that the  $f'_2(1525)$  is produced by a one-pion exchange production mechanism.  
<sup>39</sup> MARTIN 79 uses the PAWLICKI 77 data with different input value of the  $f'_2(1525) \rightarrow K\bar{K}$  branching ratio.

$\Gamma(\pi\pi)/\Gamma(K\bar{K})$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma_1$
<b>0.0092±0.0018 OUR FIT</b>					
<b>0.075 ±0.035</b>		AUGUSTIN	87	DM2 $J/\psi \rightarrow \gamma \pi^+ \pi^-$	
<b><math>[\Gamma(K\bar{K}^*(892) + \text{c.c.}) + \Gamma(\pi K\bar{K})]/\Gamma(K\bar{K})</math> (<math>\Gamma_4 + \Gamma_5</math>)/<math>\Gamma_1</math></b>					
<b>0.0092±0.0018 OUR FIT</b>					
<b>0.075 ±0.035</b>		AUGUSTIN	87	DM2 $J/\psi \rightarrow \gamma \pi^+ \pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.35	95	AGUILAR...	72B	HBC 3.9,4,6 $K^- p$	
<0.4	67	AMMAR	67	HBC	

$\Gamma(\pi\pi\eta)/\Gamma(K\bar{K})$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma_1$
<b>0.0092±0.0018 OUR FIT</b>					
<b>0.075 ±0.035</b>		AUGUSTIN	87	DM2 $J/\psi \rightarrow \gamma \pi^+ \pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.41	95	AGUILAR...	72B	HBC 3.9,4,6 $K^- p$	
<0.3	67	AMMAR	67	HBC	

$\Gamma(\pi^+ \pi^+ \pi^- \pi^-)/\Gamma(K\bar{K})$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma_1$
<b>0.0092±0.0018 OUR FIT</b>					
<b>0.075 ±0.035</b>		AUGUSTIN	87	DM2 $J/\psi \rightarrow \gamma \pi^+ \pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.32	95	AGUILAR...	72B	HBC 3.9,4,6 $K^- p$	

 $f'_2(1525)$  REFERENCES

DOBBS	15	PR D91 052006	S. Dobbs et al.	(NWES)
AJU	13AN	PR D87 072004	R. Aaij et al.	(LHCb Collab.)
ABLIKIM	13N	PR D87 092009	Ablikim M. et al.	(BES III Collab.)
UEHARA	13	PTEP 2013 123C01	S. Uehara et al.	(BES Collab.)
UEHARA	10A	PR D82 114031	S. Uehara et al.	(BES Collab.)
ANISOVICH	09	JMP A24 2481	V.Y. Anisovich, A.V. Sarantsev	(ZEUS Collab.)
CHEKANOV	08	PRL 101 112003	S. Chekanov et al.	(GAMS Collab.)
BINON	07	PAN 70 1713	F. Binon et al.	(GAMS Collab.)
		Translated from YAF 70 1758		
AMSLER	06	PL B639 165	C. Amisler et al.	(CBAR Collab.)
SCHEGELSKY	06A	EPJ A27 207	V.A. Schegelsky et al.	(CBAR Collab.)
ABLIKIM	05	PL B607 243	M. Ablikim et al.	(BES Collab.)
ABE	04	EPJ C32 323	K. ABE et al.	(BES Collab.)
CHEKANOV	04	PL B578 33	S. Chekanov et al.	(ZEUS Collab.)
BAI	03G	PR D68 052003	J.Z. Bai et al.	(BES Collab.)
TIKHOMIROV	03	PAN 66 828	G.D. Tikhomirov et al.	(BES Collab.)
		Translated from YAF 66 860		
AMSLER	02	EPJ C23 29	C. Amisler et al.	(L3 Collab.)
ACCIARRI	01H	PL B501 173	M. Acciarri et al.	(WA 102 Collab.)
BARBERIS	00E	PL B479 59	D. Barberis et al.	(Omega Expt.)
BARBERIS	99	PL B453 305	D. Barberis et al.	(Omega Expt.)
BARKOV	99	JETPL 70 248	B.P. Barkov et al.	(Omega Expt.)
		Translated from ZETFP 70 242		
ABREU	96C	PL B379 309	P. Abreu et al.	(DELPHI Collab.)
BAI	96C	PRL 77 3959	J.Z. Bai et al.	(BES Collab.)
ACCIARRI	95J	PL B363 118	M. Acciarri et al.	(L3 Collab.)
PROKOSHIN	91	SPD 36 155	Y.D. Prokoshin	(GAM2, GAM4 Collab.)
		Translated from DANS 316 900		
ALBRECHT	90G	ZPHY C48 183	H. Albrecht et al.	(ARGUS Collab.)
PDG	90	PL B239 1	J.J. Hernandez et al.	(IFIC, BOST, CIT+)
BEHREND	89C	ZPHY C43 91	H.J. Behrend et al.	(CELLO Collab.)
ASTON	88D	NP B301 525	D. Aston et al.	(SLAC, NAGO, CINC, INUS)
AUGUSTIN	88	PRL 60 2238	J.E. Augustin et al.	(DM2 Collab.)
BERGER	88	ZPHY C37 329	C. Berger et al.	(PLUTO Collab.)
FALVARD	88	PR D38 2706	A. Falvard et al.	(CLER, FRAS, LALO+)
AUGUSTIN	87	ZPHY C36 369	J.E. Augustin et al.	(LALO, CLER, FRAS+)
BALTRUSAITIS	87	PR D35 2077	R.M. Baltrusaitis et al.	(Mark III Collab.)
AIHARA	86B	PRL 57 404	H. Aihara et al.	(TTC-2γ Collab.)
BOLONKIN	86	SJNP 43 776	B.V. Bolonkin et al.	(ITEP) JP
		Translated from YAF 43 1211		
LONGACRE	86	PL B177 223	R.S. Longacre et al.	(BNL, BRAN, CUNY+)
ALTHOFF	83	PL D121 216	M. Althoff et al.	(TASSO Collab.)
ARMSTRONG	83B	NP B224 193	T.A. Armstrong et al.	(BARI, BIRM, CERN+)
AGUILAR...	81B	ZPHY C8 313	M. Aguilar-Benitez et al.	(CERN, COEF+)
ALHARRAN	81	NP B191 26	S. Alharran et al.	(BIRM, CERN, GLAS+)
CHABAUD	81	APP B12 575	V. Chabaud et al.	(CERN, CRAC, MPIM)
COSTA...	80	NP B175 402	G. Costa de Beauregard et al.	(BARI, BONN+)
GORLICH	80	NP B174 16	L. Gorlich et al.	(CRAC, MPIM, CERN+)
CORDEN	79	NP B157 250	M.J. Corden et al.	(BIRM, RHEL, TELA+)
MARTIN	79	NP B158 520	A.D. Martin, E.N. Ozmutlu	(DURH)
POLYCHRO...	79	PR D19 1317	V.A. Polychronakos et al.	(NDAM, ANL)
BARREIRO	77	NP B121 237	F. Barreiro et al.	(CERN, AMST, NIJ+)
EVANGELIS...	77	NP B127 384	C. Evangelista et al.	(BARI, BONN, CERN+)
PAWLICKI	77	PR D15 3196	A.J. Pawlicki et al.	(ANL) IJP
BRANDENB...	76C	NP B104 413	W.W. Brandenbourg et al.	(SLAC)
BEUSCH	75B	PL 60B 101	G.W. Beusch et al.	(CERN, ETH)
AGUILAR...	72B	PR D6 29	M. Aguilar-Benitez et al.	(BNL)
AMMAR	67	PRL 19 1071	R. Ammar et al.	(NWES, ANL) JP
BARNES	67	PRL 19 964	V.E. Barnes et al.	(BNL, SYRA) IJPC
CRENNELL	66	PRL 16 1025	D.J. Crennell et al.	(BNL)

 $f_2(1565)$ 

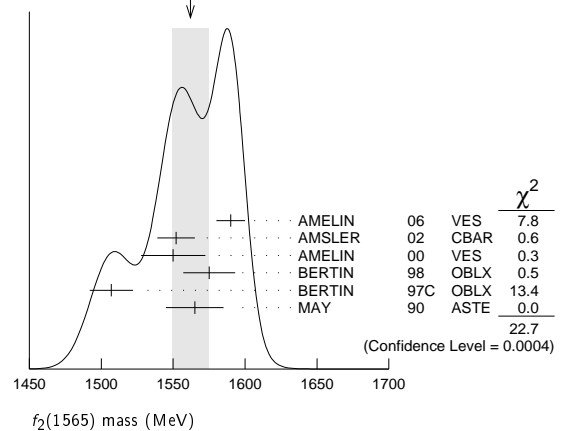
$$I^G(J^{PC}) = 0^+(2^{++})$$

OMITTED FROM SUMMARY TABLE

Seen mostly in antinucleon-nucleon annihilation. Needs confirmation in other channels.

 $f_2(1565)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1562±13 OUR AVERAGE</b>			Error includes scale factor of 2.1. See the ideogram below.
1590±10	<sup>1</sup> AMELIN	06	VES 36 $\pi^- p \rightarrow \omega \omega n$
1552±13	<sup>2</sup> AMSLER	02	CBAR 0.9 $\bar{p} p \rightarrow \pi^0 \eta \eta, \pi^0 \pi^0 \pi^0$
1550±10±20	AMELIN	00	VES 37 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
1575±18	BERTIN	98	OBLX 0.05-0.405 $\bar{p} p \rightarrow \pi^+ \pi^+ \pi^-$
1507±15	<sup>2</sup> BERTIN	97C	OBLX 0.0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
1565±20	MAY	90	ASTE 0.0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1560±15	<sup>3</sup> ANISOVICH	09	RVUE 0.0 $\bar{p} p, \pi N$
1598±11± 9	BAKER	99B	SPEC 0 $\bar{p} p \rightarrow \omega \omega \pi^0$
1534±20	<sup>4</sup> ABELE	96C	RVUE Compilation
~1552	<sup>5</sup> AMSLER	95D	CBAR 0.0 $\bar{p} p \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta \eta, \pi^0 \pi^0 \eta$
1598±72	BALOSHIN	95	SPEC 40 $\pi^- C \rightarrow K_S^0 K_S^0 X$
1566 $\pm^{+80}_{-50}$	<sup>6</sup> ANISOVICH	94	CBAR 0.0 $\bar{p} p \rightarrow 3\pi^0, \eta \eta \pi^0$
1502± 9	ADAMO	93	OBLX $\bar{p} p \rightarrow \pi^+ \pi^+ \pi^-$
1488±10	<sup>7</sup> ARMSTRONG	93C	E760 $\bar{p} p \rightarrow \pi^0 \eta \eta \rightarrow 6\gamma$
1508±10	<sup>7</sup> ARMSTRONG	93D	E760 $\bar{p} p \rightarrow 3\pi^0 \rightarrow 6\gamma$
1525±10	<sup>7</sup> ARMSTRONG	93D	E760 $\bar{p} p \rightarrow \eta \pi^0 \pi^0 \rightarrow 6\gamma$
~1504	<sup>8</sup> WEIDENAUER	93A	ASTE 0.0 $\bar{p} N \rightarrow 3\pi^- 2\pi^+$
1540±15	<sup>7</sup> ADAMO	92	OBLX $\bar{p} p \rightarrow \pi^+ \pi^+ \pi^-$
1515±10	<sup>9</sup> AKER	91	CBAR 0.0 $\bar{p} p \rightarrow 3\pi^0$
1477± 5	BRIDGES	86C	DBC 0.0 $\bar{p} N \rightarrow 3\pi^- 2\pi^+$

<sup>1</sup> Supersedes the  $\omega\omega$  state of BELADIDZE 92B earlier assigned to the  $f_2(1640)$ .<sup>2</sup> T-matrix pole.<sup>3</sup> On sheet II in a two-pole solution.<sup>4</sup> T-matrix pole, large coupling to  $\rho\rho$  and  $\omega\omega$ , could be  $f_2(1640)$ .<sup>5</sup> Coupled-channel analysis of AMSLER 95B, AMSLER 95C, and AMSLER 94D.<sup>6</sup> From a simultaneous analysis of the annihilations  $\bar{p} p \rightarrow 3\pi^0, \pi^0 \eta \eta$  including AKER 91 data.<sup>7</sup>  $J^P$  not determined, could be partly  $f_0(1500)$ .<sup>8</sup>  $J^P$  not determined.<sup>9</sup> Superseded by AMSLER 95B.WEIGHTED AVERAGE  
1562±13 (Error scaled by 2.1) $f_2(1565)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>134± 8 OUR AVERAGE</b>			
140± 11	<sup>10</sup> AMELIN	06	VES 36 $\pi^- p \rightarrow \omega \omega n$
113± 23	<sup>11</sup> AMSLER	02	CBAR 0.9 $\bar{p} p \rightarrow \pi^0 \eta \eta, \pi^0 \pi^0 \pi^0$
130± 20±40	AMELIN	00	VES 37 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
119± 24	BERTIN	98	OBLX 0.05-0.405 $\bar{p} p \rightarrow \pi^+ \pi^+ \pi^-$
130± 20	<sup>11</sup> BERTIN	97C	OBLX 0.0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
170± 40	MAY	90	ASTE 0.0 $\bar{p} p \rightarrow \pi^+ \pi^- \pi^0$



See key on page 601

## Meson Particle Listings

 $\rho(1570)$ ,  $h_1(1595)$ ,  $\pi_1(1600)$  $\rho(1570)$  BRANCHING RATIOS

$\Gamma(\phi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE				
not seen	ABELE	97H	CBAR $\bar{p}p \rightarrow K_L^0 K_S^0 \pi^0 \pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.01	7 DONNACHIE	91	RVUE	
7 Using data from BISELLO 91B, DOLINSKY 86, and ALBRECHT 87L.				

$\Gamma(\phi\pi)/\Gamma(\omega\pi)$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_3$
VALUE					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>0.5	95	BITYUKOV	87	SPEC $32.5 \pi^- p \rightarrow \phi \pi^0 n$	

 $\rho(1570)$  REFERENCES

AUBERT	08S	PR D77 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABELE	97H	PL B415 280	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BISELLO	91B	NPBPS B21 111	D. Bisello	(DM2 Collab.)
DONNACHIE	91	ZPHY C51 689	A. Donnachie, A.B. Clegg	(MCHS, LANC)
ALBRECHT	87L	PL B185 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AULCHENKO	87B	JETPL 45 145	V.M. Aulchenko <i>et al.</i>	(NOVO)
		Translated from ZETFP 45 118		
BITYUKOV	87	PL B188 383	S.I. Bityukov <i>et al.</i>	(SERP)
DOLINSKY	86	PL B174 453	S.I. Dolinsky <i>et al.</i>	(NOVO)

 **$h_1(1595)$** 

$$I^G(J^{PC}) = 0^-(1^{+-})$$

OMITTED FROM SUMMARY TABLE

Seen in a partial-wave analysis of the  $\omega\eta$  system produced in the reaction  $\pi^- p \rightarrow \omega\eta n$  at 18 GeV/c. $h_1(1595)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1594 ± 15 <sup>+10</sup> <sub>-60</sub>	EUGENIO	01	SPEC $18 \pi^- p \rightarrow \omega\eta n$

 $h_1(1595)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
384 ± 60 <sup>+70</sup> <sub>-100</sub>	EUGENIO	01	SPEC $18 \pi^- p \rightarrow \omega\eta n$

 $h_1(1595)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\omega\eta$	seen

 $h_1(1595)$  REFERENCES

EUGENIO	01	PL B497 190	P. Eugenio <i>et al.</i>
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 **$\pi_1(1600)$** 

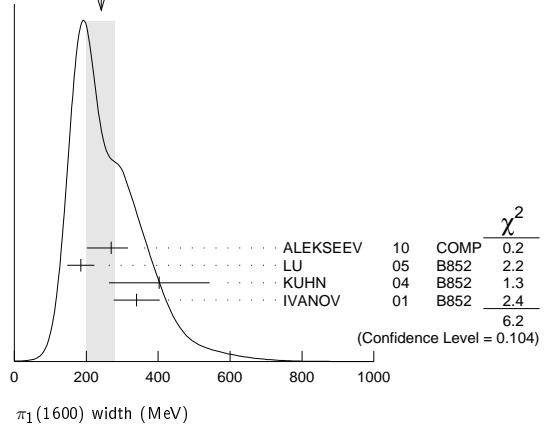
$$I^G(J^{PC}) = 1^-(1^{-+})$$

 $\pi_1(1600)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
1662 <sup>+8</sup> <sub>-9</sub> OUR AVERAGE				
1660 ± 10 <sup>+0</sup> <sub>-64</sub>	420k	ALEKSEEV	10	COMP $190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
1664 ± 8 ± 10	145k	1 LU	05	B852 $18 \pi^- p \rightarrow \omega \pi^- \pi^0 p$
1709 ± 24 ± 41	69k	2 KUHN	04	B852 $18 \pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$
1597 ± 10 <sup>+45</sup> <sub>-10</sub>		2 IVANOV	01	B852 $18 \pi^- p \rightarrow \eta' \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1593 ± 8 <sup>+29</sup> <sub>-47</sub>	2,3 ADAMS	98B	B852	$18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$

<sup>1</sup> May be a different state: natural and unnatural parity exchanges.<sup>2</sup> Natural parity exchange.<sup>3</sup> Superseded by DZIERBA 06 excluding this state in a more refined PWA analysis, with 2.6 M events of  $\pi^- p \rightarrow \pi^- \pi^- \pi^+ p$  and 3 M events of  $\pi^- p \rightarrow \pi^- \pi^0 \pi^0 p$  of E852 data. $\pi_1(1600)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
241 ± 40 OUR AVERAGE				Error includes scale factor of 1.4. See the ideogram below.
269 ± 21 <sup>+42</sup> <sub>-64</sub>	420k	ALEKSEEV	10	COMP $190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
185 ± 25 ± 28	145k	4 LU	05	B852 $18 \pi^- p \rightarrow \omega \pi^- \pi^0 p$
403 ± 80 ± 115	69k	5 KUHN	04	B852 $18 \pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$
340 ± 40 ± 50		5 IVANOV	01	B852 $18 \pi^- p \rightarrow \eta' \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
168 ± 20 <sup>+150</sup> <sub>-12</sub>	5,6 ADAMS	98B	B852	$18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$

<sup>4</sup> May be a different state: natural and unnatural parity exchanges.<sup>5</sup> Natural parity exchange.<sup>6</sup> Superseded by DZIERBA 06 excluding this state in a more refined PWA analysis, with 2.6 M events of  $\pi^- p \rightarrow \pi^- \pi^- \pi^+ p$  and 3 M events of  $\pi^- p \rightarrow \pi^- \pi^0 \pi^0 p$  of E852 data.WEIGHTED AVERAGE  
241 ± 40 (Error scaled by 1.4) $\pi_1(1600)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi\pi\pi$	not seen
$\Gamma_2$ $\rho^0 \pi^-$	not seen
$\Gamma_3$ $f_2(1270) \pi^-$	not seen
$\Gamma_4$ $b_1(1235) \pi$	seen
$\Gamma_5$ $\eta'(958) \pi^-$	seen
$\Gamma_6$ $f_1(1285) \pi$	seen

 $\pi_1(1600)$  BRANCHING RATIOS

$\Gamma(\rho^0 \pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE				
not seen	NOZAR	09	CLAS $\gamma p \rightarrow 2\pi^+ \pi^- n$	
not seen	7 DZIERBA	06	B852 $18 \pi^- p$	

<sup>7</sup> From the PWA analysis of 2.6 M  $\pi^- p \rightarrow \pi^- \pi^- \pi^+ p$  and 3 M events of  $\pi^- p \rightarrow \pi^- \pi^0 \pi^0 p$  of E852 data. Supersedes ADAMS 98B.

$\Gamma(f_2(1270) \pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
VALUE				
not seen	8 DZIERBA	06	B852 $18 \pi^- p$	

<sup>8</sup> From the PWA analysis of 2.6 M  $\pi^- p \rightarrow \pi^- \pi^- \pi^+ p$  and 3 M events of  $\pi^- p \rightarrow \pi^- \pi^0 \pi^0 p$  of E852 data. Supersedes CHUNG 02.

$\Gamma(b_1(1235) \pi)/\Gamma_{\text{total}}$	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE					
seen	35280	9 BAKER	03	SPEC $\bar{p}p \rightarrow \omega \pi^+ \pi^- \pi^0$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

seen	145k	LU	05	B852 $18 \pi^- p \rightarrow \omega \pi^- \pi^0 p$	
------	------	----	----	--	--

<sup>9</sup>  $B((b_1 \pi)_{D\text{-wave}})/B((b_1 \pi)_{S\text{-wave}}) = 0.3 \pm 0.1$ .

$\Gamma(\eta'(958) \pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE				
seen	IVANOV	01	B852 $18 \pi^- p \rightarrow \eta' \pi^- p$	

$\Gamma(f_1(1285) \pi)/\Gamma(\eta'(958) \pi^-)$	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma_5$
VALUE					
3.80 ± 0.78	69k	10 KUHN	04	B852 $18 \pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$	

<sup>10</sup> Using  $\eta'(958) \pi$  data from IVANOV 01. $\pi_1(1600)$  REFERENCES

ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i>	(COMPASS Collab.)
NOZAR	09	PRL 102 102002	M. Nozar <i>et al.</i>	(JLab CLAS Collab.)
DZIERBA	06	PR D73 072001	A.R. Dzierba <i>et al.</i>	(BNL E852 Collab.)
LU	05	PRL 94 032002	M. Lu <i>et al.</i>	(BNL E852 Collab.)
KUHN	04	PL B595 109	J. Kuhn <i>et al.</i>	(BNL E852 Collab.)
BAKER	03	PL B563 140	C.A. Baker <i>et al.</i>	
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
IVANOV	01	PRL 86 3977	E.I. Ivanov <i>et al.</i>	(BNL E852 Collab.)
ADAMS	98B	PRL 81 5760	G.S. Adams <i>et al.</i>	(BNL E852 Collab.)

# Meson Particle Listings

## $a_1(1640)$ , $f_2(1640)$

### $a_1(1640)$

$I^G(J^{PC}) = 1^-(1^{++})$

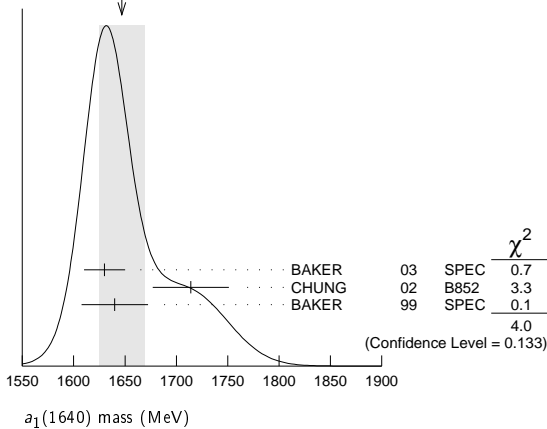
OMITTED FROM SUMMARY TABLE

Seen in the amplitude analysis of the  $3\pi^0$  system produced in  $\bar{p}p \rightarrow 4\pi^0$ . Possibly seen in the study of the hadronic structure in decay  $\tau \rightarrow 3\pi\nu_\tau$  (ABREU 98G and ASNER 00). Needs confirmation.

#### $a_1(1640)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1647±22 OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.		
1630±20	35280	<sup>1</sup> BAKER	03 SPEC	$\bar{p}p \rightarrow \omega\pi^+\pi^-\pi^0$
1714± 9±36		CHUNG	02 B852	$18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$
1640±12±30		BAKER	99 SPEC	$1.94\bar{p}p \rightarrow 4\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1670±90		BELLINI	85 SPEC	$40\pi^-A \rightarrow \pi^-\pi^+\pi^-A$

WEIGHTED AVERAGE  
1647±22 (Error scaled by 1.4)



<sup>1</sup> Using the  $a_1(1260)$  mass and width results of BOWLER 88.

#### $a_1(1640)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>254± 30 OUR AVERAGE</b>		Error includes scale factor of 1.1.		
225± 30	35280	<sup>2</sup> BAKER	03 SPEC	$\bar{p}p \rightarrow \omega\pi^+\pi^-\pi^0$
308± 37±62		CHUNG	02 B852	$18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$
300± 22±40		BAKER	99 SPEC	$1.94\bar{p}p \rightarrow 4\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
300±100		BELLINI	85 SPEC	$40\pi^-A \rightarrow \pi^-\pi^+\pi^-A$

<sup>2</sup> Using the  $a_1(1260)$  mass and width results of BOWLER 88.

#### $a_1(1640)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \pi\pi\pi$	seen
$\Gamma_2 f_2(1270)\pi$	seen
$\Gamma_3 \sigma\pi$	seen
$\Gamma_4 \rho\pi S\text{-wave}$	seen
$\Gamma_5 \rho\pi D\text{-wave}$	seen
$\Gamma_6 \omega\pi\pi$	seen
$\Gamma_7 f_1(1285)\pi$	seen
$\Gamma_8 a_1(1260)\eta$	not seen

#### $a_1(1640)$ BRANCHING RATIOS

$\Gamma(f_2(1270)\pi)/\Gamma(\sigma\pi)$	$\Gamma_2/\Gamma_3$
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.24±0.07	BAKER 99 SPEC 1.94 $\bar{p}p \rightarrow 4\pi^0$

$\Gamma(\rho\pi D\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
seen	CHUNG 02 B852 $18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$
seen	AMELIN 95B VES $36\pi^-A \rightarrow \pi^+\pi^-\pi^-A$

$\Gamma(\omega\pi\pi)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
seen	35280 <sup>3</sup> BAKER 03 SPEC $\bar{p}p \rightarrow \omega\pi^+\pi^-\pi^0$

$\Gamma(f_1(1285)\pi)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
not seen	KUHN 04 B852 $18\pi^-p \rightarrow \eta\pi^+\pi^-\pi^-p$
seen	LEE 94 MPS2 $18\pi^-p \rightarrow K^+\bar{K}^0\pi^-\pi^-p$

$\Gamma(a_1(1260)\eta)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
not seen	KUHN 04 B852 $18\pi^-p \rightarrow \eta\pi^+\pi^-\pi^-p$

<sup>3</sup> Assuming the  $\omega\rho$  mechanism for the  $\omega\pi\pi$  state.

#### $a_1(1640)$ REFERENCES

KUHN 04	PL B595 109	J. Kuhn <i>et al.</i>	(BNL E852 Collab.)
BAKER 03	PL B563 140	C.A. Baker <i>et al.</i>	
CHUNG 02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
ASNER 00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BAKER 99	PL B449 114	C.A. Baker <i>et al.</i>	
ABREU 98G	PL B426 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AMELIN 95B	PL B356 595	D.V. Amelin <i>et al.</i>	(SERP, TBIL)
LEE 94	PL B323 227	J.H. Lee <i>et al.</i>	(BNL, IND, KYUN, MASD+)
BOWLER 88	PL B209 99	M.G. Bowler	(OXF)
BELLINI 85	SJNP 41 781	D. Bellini <i>et al.</i>	
Translated from YAF 41 1223.			

### $f_2(1640)$

$I^G(J^{PC}) = 0^+(2^{++})$

OMITTED FROM SUMMARY TABLE

#### $f_2(1640)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1639± 6 OUR AVERAGE</b>		Error includes scale factor of 1.2.	
1620±16	BUGG 95	MRK3	$J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
1647± 7	ADAMO 92	OBLX	$\bar{\pi}p \rightarrow 3\pi^+2\pi^-$
1635± 7	ALDE 90	GAM2	$38\pi^-p \rightarrow \omega\omega n$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1640± 5	AMSLER 06	CBAR	$0.9\bar{p}p \rightarrow K^+K^-\pi^0$
1659± 6	VLADIMIRSK..06	SPEC	$40\pi^-p \rightarrow K_S^0 K_S^0 n$
1643± 7	<sup>1</sup> ALDE 89B	GAM2	$38\pi^-p \rightarrow \omega\omega n$

<sup>1</sup> Superseded by ALDE 90.

#### $f_2(1640)$ WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>99+60-40 OUR AVERAGE</b>		Error includes scale factor of 2.9.		
140+60-20		BUGG 95	MRK3	$J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
58±20		ADAMO 92	OBLX	$\bar{\pi}p \rightarrow 3\pi^+2\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
44± 9		AMSLER 06	CBAR	$0.9\bar{p}p \rightarrow K^+K^-\pi^0$
152±18		VLADIMIRSK..06	SPEC	$40\pi^-p \rightarrow K_S^0 K_S^0 n$
< 70	90	ALDE 90	GAM2	$38\pi^-p \rightarrow \omega\omega n$

#### $f_2(1640)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \omega\omega$	seen
$\Gamma_2 4\pi$	seen
$\Gamma_3 K\bar{K}$	seen

#### $f_2(1640)$ BRANCHING RATIOS

$\Gamma(K\bar{K})/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
seen	AMSLER 06 CBAR $0.9\bar{p}p \rightarrow K^+K^-\pi^0$

#### $f_2(1640)$ REFERENCES

AMSLER 06	PL B639 165	C. Amsler <i>et al.</i>	(CBAR Collab.)
VLADIMIRSK..06	PAN 69 493	V.V. Vladimirovsky <i>et al.</i>	(ITEP, Moscow)
Translated from YAF 69 515.			
BUGG 95	PL B353 378	D.V. Bugg <i>et al.</i>	(LOQM, PNPI, WASH)JP
ADAMO 92	PL B267 368	A. Adamo <i>et al.</i>	(OBELIX Collab.)
ALDE 90	PL B241 600	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP+)
ALDE 89B	PL B216 451	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP+)IGJPC





# Meson Particle Listings

## $\omega(1650)$ , $\omega_3(1670)$

### $\Gamma(\omega\pi\pi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_2/\Gamma \times \Gamma_4/\Gamma$

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
7.0 $\pm$ 0.5		AUBERT	07AU BABR	10.6 $e^+e^- \rightarrow \omega\pi^+\pi^-\gamma$
4.1 $\pm$ 0.9 $\pm$ 1.3	1.2M 18,19	ACHASOV	03D RVUE	0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
5.40 $\pm$ 0.95		24 AKHMETSHIN 00D	CMD2	1.2–1.38 $e^+e^- \rightarrow \omega\pi^+\pi^-$
3.18 $\pm$ 0.80		20,21 CLEGG	94 RVUE	
6.07 $\pm$ 0.61	750 22,23	ANTONELLI 92	DM2	1.34–2.4 $e^+e^- \rightarrow \rho\pi, \omega\pi\pi$

### $\Gamma(\omega\eta)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_3/\Gamma \times \Gamma_4/\Gamma$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
0.57 $\pm$ 0.06		13	AUBERT	06D BABR	10.6 $e^+e^- \rightarrow \omega\eta\gamma$
<6	90	25	AKHMETSHIN 03B	CMD2	$e^+e^- \rightarrow \eta\pi^0\gamma$
17 From a fit with contributions from $\omega(782)$ , $\phi(1020)$ , $\omega(1420)$ , and $\omega(1650)$ .					
18 Calculated by us from the cross section at the peak.					
19 From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the $\pi^+\pi^-\pi^0$ and ANTONELLI 92 on the $\omega\pi^+\pi^-$ final states. Supersedes ACHASOV 99E and ACHASOV 02E.					
20 From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.					
21 From the partial and leptonic width given by the authors.					
22 From the combined fit of the $\rho\pi$ and $\omega\pi\pi$ final states.					
23 From the product of the leptonic width and partial branching ratio given by the authors.					
24 Using the data of AKHMETSHIN 00D and ANTONELLI 92. The $\rho\pi$ dominance for the energy dependence of the $\omega(1420)$ and $\omega(1650)$ width assumed.					
25 $\omega(1650)$ mass and width fixed at 1700 MeV and 250 MeV, respectively.					

### $\omega(1650)$ BRANCHING RATIOS

### $\Gamma(\omega\pi\pi)/\Gamma_{\text{total}} \qquad \Gamma_2/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$\sim$ 0.35	1.2M	26	ACHASOV	03D RVUE 0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.620 $\pm$ 0.014		27	HENNER	02 RVUE 1.2–2.0 $e^+e^- \rightarrow \rho\pi, \omega\pi\pi$

### $\Gamma(\rho\pi)/\Gamma_{\text{total}} \qquad \Gamma_1/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$\sim$ 0.65	1.2M	26	ACHASOV	03D RVUE 0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
0.380 $\pm$ 0.014		27	HENNER	02 RVUE 1.2–2.0 $e^+e^- \rightarrow \rho\pi, \omega\pi\pi$

### $\Gamma(e^+e^-)/\Gamma_{\text{total}} \qquad \Gamma_4/\Gamma$

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$\sim$ 18	1.2M	27,28	ACHASOV	03D RVUE 0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
32 $\pm$ 1		27	HENNER	02 RVUE 1.2–2.0 $e^+e^- \rightarrow \rho\pi, \omega\pi\pi$

26 From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the  $\pi^+\pi^-\pi^0$  and ANTONELLI 92 on the  $\omega\pi^+\pi^-$  final states. Supersedes ACHASOV 99E and ACHASOV 02E.

27 Assuming that the  $\omega(1650)$  decays into  $\rho\pi$  and  $\omega\pi\pi$  only.

28 Calculated by us from the cross section at the peak.

### $\omega(1650)$ REFERENCES

AULCHENKO	15A	JETP 121 27 Translated from ZETF 148 34.	V.M. Aulchenko <i>et al.</i>	(SND Collab.)
AUBERT	07AU	PR D76 092005	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	06D	PR D73 052003	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	04N	PR D70 072004	B. Aubert <i>et al.</i>	(BABAR Collab.)
ACHASOV	03D	PR D68 052006	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	03B	PL B562 173	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ACHASOV	02E	PR D66 032001	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
HENNER	02	EPJ C26 3	V.K. Henner <i>et al.</i>	
ACHASOV	01E	PR D63 072002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
EUGENIO	01	PL B497 190	P. Eugenio <i>et al.</i>	
AKHMETSHIN	00D	PL B489 125	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ACHASOV	99E	PL B462 365	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	98H	PR D57 4334	N.N. Achasov, A.A. Kozhevnikov	
CLEGG	94	ZPHY C62 455	A.B. Clegg, A. Donnachie	(LANC, MCHS)
ANTONELLI	92	ZPHY C56 15	A. Antonelli <i>et al.</i>	(DM2 Collab.)
BISELLO	91C	ZPHY C52 227	D. Bisello <i>et al.</i>	(DM2 Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
BISELLO	88B	ZPHY C39 13	D. Bisello <i>et al.</i>	(PADO, CLER, FRAS+)
BARKOV	87	JETPL 46 164 Translated from ZETFP 46 132	L.M. Barkov <i>et al.</i>	(NOVO)
ATKINSON	83B	PL 127B 132	M. Atkinson <i>et al.</i>	(BOONN, CERN, GLAS+)
CORDIER	81	PL 106B 155	A. Cordier <i>et al.</i>	(ORSAY)
IVANOV	81	PL 107B 297	P.M. Ivanov <i>et al.</i>	(NOVO)
ESPOSITO	80	LCN 28 195	E. Esposito <i>et al.</i>	(FRAS, NAPL, PADO+)
COSME	79	NP B152 215	G. Cosme <i>et al.</i>	(IPN)

## $\omega_3(1670)$

$$I^G(J^{PC}) = 0^-(3^{--})$$

### $\omega_3(1670)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
1667 $\pm$ 4	OUR AVERAGE			
1665.3 $\pm$ 5.2 $\pm$ 4.5	23400	AMELIN	96 VES	36 $\pi^-\rho \rightarrow \pi^+\pi^-\pi^0\eta$
1685 $\pm$ 20	60	BAUBILLIER	79 HBC	8.2 $K^-\rho$ backward
1673 $\pm$ 12	430	1,2 BALTAY	78E HBC	15 $\pi^+\rho \rightarrow \Delta 3\pi$
1650 $\pm$ 12		CORDEN	78B OMEG	8–12 $\pi^-\rho \rightarrow N3\pi$
1669 $\pm$ 11	600	2 WAGNER	75 HBC	7 $\pi^+\rho \rightarrow \Delta^{++}3\pi$
1678 $\pm$ 14	500	DIAZ	74 DBC	6 $\pi^+n \rightarrow p3\pi^0$
1660 $\pm$ 13	200	DIAZ	74 DBC	6 $\pi^+n \rightarrow \rho\omega\pi^0\pi^0$
1679 $\pm$ 17	200	MATTHEWS	71D DBC	7.0 $\pi^+n \rightarrow p3\pi^0$
1670 $\pm$ 20		KENYON	69 DBC	8 $\pi^+n \rightarrow p3\pi^0$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$\sim$ 1700	110	1 CERRADA	77B HBC	4.2 $K^-\rho \rightarrow \Lambda 3\pi$
1695 $\pm$ 20		BARNES	69B HBC	4.6 $K^-\rho \rightarrow \omega 2\pi X$
1636 $\pm$ 20		ARMENISE	68B DBC	5.1 $\pi^+n \rightarrow p3\pi^0$
1 Phase rotation seen for $J^P = 3^--\rho\pi$ wave.				
2 From a fit to $I(J^P) = 0(3^{--})\rho\pi$ partial wave.				

### $\omega_3(1670)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
168 $\pm$ 10	OUR AVERAGE			
149 $\pm$ 19 $\pm$ 7	23400	AMELIN	96 VES	36 $\pi^-\rho \rightarrow \pi^+\pi^-\pi^0\eta$
160 $\pm$ 80	60	3 BAUBILLIER	79 HBC	8.2 $K^-\rho$ backward
173 $\pm$ 16	430	4,5 BALTAY	78E HBC	15 $\pi^+\rho \rightarrow \Delta 3\pi$
253 $\pm$ 39		CORDEN	78B OMEG	8–12 $\pi^-\rho \rightarrow N3\pi$
173 $\pm$ 28	600	3,5 WAGNER	75 HBC	7 $\pi^+\rho \rightarrow \Delta^{++}3\pi$
167 $\pm$ 40	500	DIAZ	74 DBC	6 $\pi^+n \rightarrow p3\pi^0$
122 $\pm$ 39	200	DIAZ	74 DBC	6 $\pi^+n \rightarrow \rho\omega\pi^0\pi^0$
155 $\pm$ 40	200	3 MATTHEWS	71D DBC	7.0 $\pi^+n \rightarrow p3\pi^0$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
90 $\pm$ 20		BARNES	69B HBC	4.6 $K^-\rho \rightarrow \omega 2\pi$
100 $\pm$ 40		KENYON	69 DBC	8 $\pi^+n \rightarrow p3\pi^0$
112 $\pm$ 60		ARMENISE	68B DBC	5.1 $\pi^+n \rightarrow p3\pi^0$
3 Width errors enlarged by us to $4\Gamma/\sqrt{N}$ ; see the note with the $K^*(892)$ mass.				
4 Phase rotation seen for $J^P = 3^--\rho\pi$ wave.				
5 From a fit to $I(J^P) = 0(3^{--})\rho\pi$ partial wave.				

### $\omega_3(1670)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \rho\pi$	seen
$\Gamma_2 \quad \omega\pi\pi$	seen
$\Gamma_3 \quad b_1(1235)\pi$	possibly seen

### $\omega_3(1670)$ BRANCHING RATIOS

### $\Gamma(\omega\pi\pi)/\Gamma(\rho\pi) \qquad \Gamma_2/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.71 $\pm$ 0.27	100	DIAZ	74 DBC	6 $\pi^+n \rightarrow p5\pi^0$

### $\Gamma(b_1(1235)\pi)/\Gamma(\rho\pi) \qquad \Gamma_3/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
possibly seen	DIAZ	74 DBC	6 $\pi^+n \rightarrow p5\pi^0$

### $\Gamma(b_1(1235)\pi)/\Gamma(\omega\pi\pi) \qquad \Gamma_3/\Gamma_2$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
>0.75	68	BAUBILLIER	79 HBC	8.2 $K^-\rho$ backward

### $\omega_3(1670)$ REFERENCES

AMELIN	96	ZPHY C70 71	D.V. Amelin <i>et al.</i>	(SERP, TBIL)
BAUBILLIER	79	PL 89B 131	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
BALTAY	78E	PRL 40 87	C. Baltay, C.V. Cautis, M. Kalelkar	(COLU)JP
CORDEN	78B	NP B138 235	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TELA+)
CERRADA	77B	NP B126 241	M. Cerrada <i>et al.</i>	(AMST, CERN, NIJ+ )JP
WAGNER	75	PL 58B 201	F. Wagner, M. Tabak, D.M. Chew	(LBL)JP
DIAZ	74	PRL 32 260	J. Diaz <i>et al.</i>	(CASE, CMU)
MATTHEWS	71D	PR D3 2561	J.A.J. Matthews <i>et al.</i>	(TNTO, WISC)
BARNES	69B	PRL 23 142	V.E. Barnes <i>et al.</i>	(BNL)
KENYON	69	PRL 23 146	I.R. Kenyon <i>et al.</i>	(BNL, UCND, ORNL)
ARMENISE	68B	PL 26B 336	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ+)

See key on page 601

## Meson Particle Listings

 $\pi_2(1670)$  $\pi_2(1670)$ 

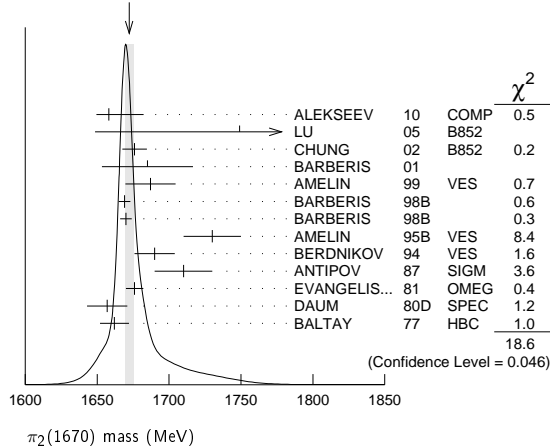
$$J^G(J^{PC}) = 1^-(2^{-+})$$

 $\pi_2(1670)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1672.2 ± 3.0 OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.			
1658 ± 3 ± $^{24}_8$	420k	ALEKSEEV	10	COMP	$190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
1749 ± 10 ± 100	145k	LU	05	B852	$18 \pi^- p \rightarrow \omega \pi^- \pi^0 p$
1676 ± 3 ± 8		1 CHUNG	02	B852	$18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
1685 ± 10 ± 30		2 BARBERIS	01		$450 \rho p \rightarrow \rho_f 3\pi^0 p_S$
1687 ± 9 ± 15		AMELIN	99	VES	$37 \pi^- A \rightarrow \omega \pi^- \pi^0 A^*$
1669 ± 4		BARBERIS	98B		$450 \rho p \rightarrow \rho_f \rho \pi p_S$
1670 ± 4		BARBERIS	98B		$450 \rho p \rightarrow \rho_f f_2(1270) \pi p_S$
1730 ± 20		3 AMELIN	95B	VES	$36 \pi^- A \rightarrow \pi^+ \pi^- \pi^- A$
1690 ± 14		4 BERDNIKOV	94	VES	$37 \pi^- A \rightarrow K^+ K^- \pi^- A$
1710 ± 20	700	ANTIPOV	87	SIGM	$50 \pi^- Cu \rightarrow \mu^+ \mu^- \pi^- Cu$
1676 ± 6		4 EVANGELIS...	81	OMEG	$12 \pi^- p \rightarrow 3\pi p$
1657 ± 14		4.5 DAUM	80D	SPEC	$63-94 \pi p \rightarrow 3\pi X$
1662 ± 10	2000	4 BALTAY	77	HBC	$15 \pi^+ p \rightarrow \rho 3\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1742 ± 31 ± 49		ANTREASYAN	90	CBAL	$e^+ e^- \rightarrow e^+ e^- \pi^0 \pi^0 \pi^0$
1624 ± 21		1 BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1622 ± 35		6 BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1693 ± 28		7 BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1710 ± 20		8 DAUM	81B	SPEC	$63.94 \pi^- p$
1660 ± 10		4 ASCOLI	73	HBC	$5-25 \pi^- p \rightarrow \rho \pi_2$

1 From  $f_2(1270) \pi$  decay.

2 From a fit to the invariant mass distribution.

3 From a fit to  $J^{PC} = 2^{-+} f_2(1270) \pi, f_0(1370) \pi$  waves.4 From a fit to  $J^P = 2^{-} S$ -wave  $f_2(1270) \pi$  partial wave.5 Clear phase rotation seen in  $2^{-} S, 2^{-} P, 2^{-} D$  waves. We quote central value and spread of single-resonance fits to three channels.6 From  $\rho \pi$  decay.7 From  $\sigma \pi$  decay.8 From a two-resonance fit to four  $2^{-} 0^{+}$  waves. This should not be averaged with all the single resonance fits.WEIGHTED AVERAGE  
1672.2±3.0 (Error scaled by 1.4) $\pi_2(1670)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>260 ± 9 ± 22</b>	420k	ALEKSEEV	10	COMP	$190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
271 ± 60 ± 250	145k	LU	05	B852	$18 \pi^- p \rightarrow \omega \pi^- \pi^0 p$
254 ± 3 ± 31		9 CHUNG	02	B852	$18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
265 ± 30 ± 40		10 BARBERIS	01		$450 \rho p \rightarrow \rho_f 3\pi^0 p_S$
168 ± 43 ± 53		AMELIN	99	VES	$37 \pi^- A \rightarrow \omega \pi^- \pi^0 A^*$
268 ± 15		BARBERIS	98B		$450 \rho p \rightarrow \rho_f \rho \pi p_S$

256 ± 15	BARBERIS	98B	450 $\rho p \rightarrow \rho_f f_2(1270) \pi p_S$
310 ± 20	11 AMELIN	95B	36 $\pi^- A \rightarrow \pi^+ \pi^- \pi^- A$
190 ± 50	12 BERDNIKOV	94	37 $\pi^- A \rightarrow K^+ K^- \pi^- A$
170 ± 80	700 ANTIPOV	87	50 $\pi^- Cu \rightarrow \mu^+ \mu^- \pi^- Cu$
260 ± 20	12 EVANGELIS...	81	12 $\pi^- p \rightarrow 3\pi p$
219 ± 20	12.13 DAUM	80D	63-94 $\pi p \rightarrow 3\pi X$
285 ± 60	2000 12 BALTAY	77	15 $\pi^+ p \rightarrow \rho 3\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
236 ± 49 ± 36	ANTREASYAN	90	$e^+ e^- \rightarrow e^+ e^- \pi^0 \pi^0 \pi^0$
304 ± 22	9 BELLINI	85	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
404 ± 108	14 BELLINI	85	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
330 ± 90	15 BELLINI	85	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
312 ± 50	16 DAUM	81B	63.94 $\pi^- p$
270 ± 60	12 ASCOLI	73	5-25 $\pi^- p \rightarrow \rho \pi_2$

9 From  $f_2(1270) \pi$  decay.

10 From a fit to the invariant mass distribution.

11 From a fit to  $J^{PC} = 2^{-+} f_2(1270) \pi, f_0(1370) \pi$  waves.12 From a fit to  $J^P = 2^{-} f_2(1270) \pi$  partial wave.13 Clear phase rotation seen in  $2^{-} S, 2^{-} P, 2^{-} D$  waves. We quote central value and spread of single-resonance fits to three channels.14 From  $\rho \pi$  decay.15 From  $\sigma \pi$  decay.16 From a two-resonance fit to four  $2^{-} 0^{+}$  waves. This should not be averaged with all the single resonance fits. $\pi_2(1670)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $3\pi$	(95.8 ± 1.4) %	
$\Gamma_2$ $\pi^+ \pi^- \pi^0$		
$\Gamma_3$ $\pi^0 \pi^0 \pi^0$		
$\Gamma_4$ $f_2(1270) \pi$	(56.3 ± 3.2) %	
$\Gamma_5$ $\rho \pi$	(31 ± 4) %	
$\Gamma_6$ $\sigma \pi$	(10.9 ± 3.4) %	
$\Gamma_7$ $\pi(\pi\pi)S$ -wave	( 8.7 ± 3.4) %	
$\Gamma_8$ $KK^*(892) + c.c.$	( 4.2 ± 1.4) %	
$\Gamma_9$ $\omega \rho$	( 2.7 ± 1.1) %	
$\Gamma_{10}$ $\pi^\pm \gamma$	( 7.0 ± 1.1) × 10 <sup>-4</sup>	
$\Gamma_{11}$ $\gamma \gamma$	< 2.8 × 10 <sup>-7</sup>	90%
$\Gamma_{12}$ $\eta \pi$		
$\Gamma_{13}$ $\pi^\pm 2\pi^+ 2\pi^-$		
$\Gamma_{14}$ $\rho(1450) \pi$	< 3.6 × 10 <sup>-3</sup>	97.7%
$\Gamma_{15}$ $b_1(1235) \pi$	< 1.9 × 10 <sup>-3</sup>	97.7%
$\Gamma_{16}$ $\eta 3\pi$		
$\Gamma_{17}$ $f_1(1285) \pi$	possibly seen	
$\Gamma_{18}$ $a_2(1320) \pi$	not seen	

## CONSTRAINED FIT INFORMATION

An overall fit to 4 branching ratios uses 6 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 1.9$  for 3 degrees of freedom.

The following off-diagonal array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_5$	-53		
$x_7$	-29	-59	
$x_8$	-8	-21	-9
$x_4$	$x_5$	$x_7$	

 $\pi_2(1670)$  PARTIAL WIDTHS

VALUE (keV)	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{10}$
<b>181 ± 11 ± 27</b>	17 ADOLPH	14	COMP	$190 \pi^- Pb \rightarrow \pi^+ \pi^- \pi^- Pb'$	
17 Primakoff reaction. Assumes incoherent $f_2(1270) \pi$ contribution to $3\pi$ final state and uses $B(\pi_2(1670) \rightarrow f_2 \pi) = 56\%$ .					

# Meson Particle Listings

## $\pi_2(1670)$ , $\phi(1680)$

$\Gamma(\gamma\gamma)$						$\Gamma_{11}$
VALUE (keV)	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<b>&lt;0.072</b>	90	18 ACCIARRI	97T	L3	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$
<0.19	90	18 ALBRECHT	97B	ARG	$e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	
1.41 $\pm 0.23 \pm 0.28$		ANTREASIAN	90	CBAL	0 $e^+e^- \rightarrow e^+e^-\pi^0\pi^0\pi^0$	
0.8 $\pm 0.3 \pm 0.12$		19 BEHREND	90c	CELL	0 $e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	
1.3 $\pm 0.3 \pm 0.2$		20 BEHREND	90c	CELL	0 $e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	
18 Decaying into $f_2(1270)\pi$ and $\rho\pi$ .						
19 Constructive interference between $f_2(1270)\pi, \rho\pi$ and background.						
20 Incoherent Ansatz.						

$\pi_2(1670) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$						$\Gamma_{211}/\Gamma$
VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;0.1</b>	95	21 SCHEGELSKY	06	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$	
21 From analysis of L3 data at 183–209 GeV.						

$\pi_2(1670)$ BRANCHING RATIOS					
$\Gamma(3\pi)/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma = (\Gamma_4+\Gamma_5+\Gamma_7)/\Gamma$		
VALUE	DOCUMENT ID				
<b>0.958±0.014 OUR FIT</b>					
$\Gamma(\pi^0\pi^0\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$			$\Gamma_3/\Gamma_2$		
VALUE	DOCUMENT ID	COMMENT			
<b>0.29±0.03±0.05</b>	22 BARBERIS	01	450	$p\bar{p} \rightarrow p_f 3\pi^0 p_s$	

$\Gamma(\rho\pi)/0.565\Gamma(f_2(1270)\pi)$				$\Gamma_5/0.565\Gamma_4$	
(With $f_2(1270) \rightarrow \pi^+\pi^-$ .)					
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.97<math>\pm</math>0.09 OUR AVERAGE</b>	Error includes scale factor of 1.9.				
0.76 $\pm$ 0.07 $\pm$ 0.10	CHUNG	02	B852	18.3 $\pi^-p \rightarrow \pi^+\pi^-\pi^-p$	
1.01 $\pm$ 0.05	BARBERIS	98B		450 $p\bar{p} \rightarrow p_f\pi^+\pi^-\pi^0p_s$	

$\Gamma(\sigma\pi)/\Gamma(f_2(1270)\pi)$						$\Gamma_6/\Gamma_4$
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>0.19<math>\pm</math>0.06 OUR AVERAGE</b>						
0.17 $\pm$ 0.02 $\pm$ 0.07	CHUNG	02	B852	18.3	$\pi^-\bar{p} \rightarrow \pi^+\pi^-\pi^-\bar{p}$	
0.24 $\pm$ 0.10	23,24 BAKER	99	SPEC	1.94	$\bar{p}p \rightarrow 4\pi^0$	

$\frac{1}{2}\Gamma(\rho\pi)/\Gamma(\pi^\pm\pi^+\pi^-)$						$\frac{1}{2}\Gamma_5/(0.565\Gamma_4+\frac{1}{2}\Gamma_5+0.624\Gamma_7)$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.29<math>\pm</math>0.04 OUR FIT</b>						
<b>0.29<math>\pm</math>0.05</b>	25 DAUM	81B	SPEC		63,94 $\pi^-\bar{p}$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$						
<0.3	BARTSCH	68	HBC	+	8 $\pi^+\bar{p} \rightarrow 3\pi p$	

$0.565\Gamma(f_2(1270)\pi)/\Gamma(\pi^\pm\pi^+\pi^-)$			$0.565\Gamma_4/(0.565\Gamma_4+\frac{1}{2}\Gamma_5+0.624\Gamma_7)$		
(With $f_2(1270) \rightarrow \pi^+\pi^-$ .)					
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>0.604<math>\pm</math>0.035 OUR FIT</b>					
<b>0.60 <math>\pm</math>0.05 OUR AVERAGE</b>	Error includes scale factor of 1.3.				
0.61 $\pm$ 0.04	25 DAUM	81B	SPEC		63,94 $\pi^-p$
0.76 $\begin{smallmatrix} +0.24 \\ -0.34 \end{smallmatrix}$	ARMENISE	69	DBC	+	5.1 $\pi^+d \rightarrow d 3\pi$
0.35 $\pm$ 0.20	BALTAY	68	HBC	+	7–8.5 $\pi^+\pi^+$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
0.59	BARTSCH	68	HBC	+	8 $\pi^+p \rightarrow 3\pi p$

$0.624\Gamma(\pi(\pi\pi)_{S\text{-wave}})/\Gamma(\pi^\pm\pi^+\pi^-)$				$0.624\Gamma_7/(0.565\Gamma_4+\frac{1}{2}\Gamma_5+0.624\Gamma_7)$	
(With $(\pi\pi)_{S\text{-wave}} \rightarrow \pi^+\pi^-$ .)					
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.10<math>\pm</math>0.04 OUR FIT</b>					
<b>0.10<math>\pm</math>0.05</b>	25 DAUM	81B SPEC	63,94 $\pi^-p$		

$\Gamma(K\bar{K}^*(892)+\text{c.c.})/\Gamma(f_2(1270)\pi)$						$\Gamma_8/\Gamma_4$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.075<math>\pm</math>0.025 OUR FIT</b>						
<b>0.075<math>\pm</math>0.025</b>	26 ARMSTRONG	82B	OMEG	—	16 $\pi^-\bar{p} \rightarrow K^+K^-\pi^-\bar{p}$	

$\Gamma(\omega\rho)/\Gamma_{\text{total}}$						$\Gamma_9/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>0.027<math>\pm</math>0.004<math>\pm</math>0.010</b>	27 AMELIN	99	VES	37	$\pi^-A \rightarrow \omega\pi^-\pi^0A^*$	

$\Gamma(\eta\pi)/\Gamma(\pi^\pm\pi^+\pi^-)$ (All $\eta$ decays.)				$\Gamma_{12}/(0.565\Gamma_4+\frac{1}{2}\Gamma_5+0.624\Gamma_7)$	
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>&lt;0.09</b>	BALTAY	68	HBC	+	$7-8.5\pi^+p$
••• We do not use the following data for averages, fits, limits, etc. •••					
<0.10	CRENNELL	70	HBC	—	$6\pi^-p \rightarrow f_2\pi^-N$

$\Gamma(\pi^\pm 2\pi^+ 2\pi^-)/\Gamma(\pi^\pm\pi^+\pi^-)$						$\Gamma_{13}/(0.565\Gamma_4+\frac{1}{2}\Gamma_5+0.624\Gamma_7)$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT		
<b>&lt;0.10</b>	CRENNELL	70	HBC	—	6 $\pi^-\bar{p} \rightarrow f_2\pi^-N$	
<0.1	BALTAY	68	HBC	+	7,8.5 $\pi^+\pi^+\bar{p}$	

$\Gamma(\rho(1450)\pi)/\Gamma_{\text{total}}$						$\Gamma_{14}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;0.0036</b>	97.7	AMELIN	99	VES	37 $\pi^-A \rightarrow \omega\pi^-\pi^0A^*$	

$\Gamma(b_1(1235)\pi)/\Gamma_{\text{total}}$						$\Gamma_{15}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;0.0019</b>	97.7	AMELIN	99	VES	37 $\pi^-A \rightarrow \omega\pi^-\pi^0A^*$	

$\Gamma(f_1(1285)\pi)/\Gamma_{\text{total}}$						$\Gamma_{17}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>possibly seen</b>	69k	KUHN	04	B852	18 $\pi^-\bar{p} \rightarrow \eta\pi^+\pi^-\pi^-\bar{p}$	

$\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$						$\Gamma_{18}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>not seen</b>	69k	KUHN	04	B852	18 $\pi^-\bar{p} \rightarrow \eta\pi^+\pi^-\pi^-\bar{p}$	

$D\text{-wave}/S\text{-wave RATIO FOR } \pi_2(1670) \rightarrow f_2(1270)\pi$						$\Gamma_{19}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>–0.18<math>\pm</math>0.06</b>	23 BAKER	99	SPEC	1.94	$\bar{p}p \rightarrow 4\pi^0$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$						
0.22 $\pm$ 0.10	25 DAUM	81B	SPEC		63,94 $\pi^-\bar{p}$	

$F\text{-wave}/P\text{-wave RATIO FOR } \pi_2(1670) \rightarrow \rho\pi$						$\Gamma_{20}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>–0.72<math>\pm</math>0.07<math>\pm</math>0.14</b>	CHUNG	02	B852	18.3	$\pi^-\bar{p} \rightarrow \pi^+\pi^-\pi^-\bar{p}$	
22 Using BARBERIS 98B.						
23 Using preliminary CBAR data.						
24 With the $\sigma\pi$ in $L=2$ and the $f_2(1270)\pi$ in $L=0$ .						
25 From a two-resonance fit to four $2^-0^+$ waves.						
26 From a partial-wave analysis of $K^+K^-\pi^-$ system.						
27 Normalized to the B( $\pi_2(1670) \rightarrow f_2\pi$ ).						

$\pi_2(1670)$ REFERENCES					
ADOLPH	14	EPJ A50 79	C. Adolph <i>et al.</i>	(COMPASS Collab.)	
ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i>	(COMPASS Collab.)	
SCHEGELSKY	06	EPJ A27 199	V.A. Schegelsky <i>et al.</i>		
LU	05	PRL 94 032002	M. Lu <i>et al.</i>	(BNL E852 Collab.)	
KUHN	04	PL B595 109	J. Kuhn <i>et al.</i>	(BNL E852 Collab.)	
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)	
BARBERIS	01	PL B507 14	D. Barberis <i>et al.</i>		
AMELIN	99	PAN 62 445	D.V. Amelin <i>et al.</i>	(VES Collab.)	
BAKER	99	PL B449 114	C.A. Baker <i>et al.</i>		
BARBERIS	98B	PL B422 399	D. Barberis <i>et al.</i>	(WA 102 Collab.)	
ACCIARRI	97T	PL B413 147	M. Acciarri <i>et al.</i>	(L3 Collab.)	
ALBRECHT	97B	ZPHY C74 469	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	
AMELIN	95B	PL B356 595	D.W. Amelin <i>et al.</i>	(SERP, TBIL)	
BERDNIKOV	94	PL B337 219	E.B. Berdnikov <i>et al.</i>	(SERP, TBIL)	
ANTREASIAN	90	ZPHY C48 561	D. Antreasian <i>et al.</i>	(Crystal Ball Collab.)	
BEHREND	90C	ZPHY C46 583	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	
ANTIPOV	87	EPL 4 403	Y.M. Antipov <i>et al.</i>	(SERP, JINR, INRM+)	
BELLINI	85	SJNP 41 781	D. Bellini <i>et al.</i>		
ARMSTRONG	82B	NP B202 1	T.A. Armstrong, B. Baccari	(AACH3, BARI, BONN+)	
DAUM	81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)	
EVANGELISTA...	81	NP B178 197	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)	
Also	80D	NP B186 594	C. Evangelista		
DAUM	81B	PL B98 285	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)	
BALTAY	77	PRL 39 591	C. Baltay, C.V. Cautis, M. Kaleikar	(COLU)JP	
ASCOLI	73	PR D7 669	G. Ascoli	(ILL, TNT0, GENO, HAMB, MILA+)	
CRENNELL	70	PRL 24 781	D.J. Crennell <i>et al.</i>	(BNL)	
ARMENISE	69	LNC 2 501	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ)	
BALTAY	68	PRL 20 887	C. Baltay <i>et al.</i>	(COLU, ROCH, RUTG, YALE)J	
BARTSCH	68	NP B7 345	J. Bartsch <i>et al.</i>	(AACH, BERL, CERN)JP	

$\phi(1680)$					
$I(J^{PC}) = 0^-(1^--)$					

$\phi(1680)$ MASS					
$e^+e^-$ PRODUCTION					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1680<math>\pm</math>20 OUR ESTIMATE</b>					
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
1689 $\pm$ 7 $\pm$ 10	4.8k	1 SHEN	09	BELL	10.6 $e^+e^- \rightarrow K^+K^-\pi^+\pi^-\gamma$
1709 $\pm$ 20 $\pm$ 43		2 AUBERT	08B	BABR	10.6 $e^+e^- \rightarrow$ hadrons
1623 $\pm$ 20	948	3 AKHMETSHIN	03	CMD2	1.05–1.38 $e^+e^- \rightarrow K_L^0 K_S^0$
$\sim 1500$		4 ACHASOV	98H	RVUE	$e^+e^- \rightarrow \pi^+\pi^-\pi^0, \omega\pi^+\pi^-, K^+K^-$
$\sim 1900$		5 ACHASOV	98H	RVUE	$e^+e^- \rightarrow K_S^0 K^\pm\pi^\mp$
1700 $\pm$ 20		6 CLEGG	94	RVUE	$e^+e^- \rightarrow K^+K^-, K_S^0 K\pi$
1657 $\pm$ 27	367	7 BISELLO	91C	DM2	$e^+e^- \rightarrow K_S^0 K^\pm\pi^\mp$
1655 $\pm$ 17		8 BISELLO	88B	DM2	$e^+e^- \rightarrow K^+K^-$
1680 $\pm$ 10		9 BUON	82	DM1	$e^+e^- \rightarrow$ hadrons
1677 $\pm$ 12		MANE	82	DM1	$e^+e^- \rightarrow K_S^0 K\pi$

- <sup>1</sup> From a fit with two incoherent Breit-Wigners.  
<sup>2</sup> From the simultaneous fit to the  $K\bar{K}^*(892) + \text{c.c.}$  and  $\phi\eta$  data from AUBERT 08s using the results of AUBERT 07AK.  
<sup>3</sup> From the combined fit of AKHMETSHIN 03 and MANE 81 also including  $\rho$ ,  $\omega$ , and  $\phi$ . Neither isospin nor flavor structure known.  
<sup>4</sup> Using data from IVANOV 81, BARKOV 87, BISELLO 88b, DOLINSKY 91, and ANTONELLI 92.  
<sup>5</sup> Using the data from BISELLO 91c.  
<sup>6</sup> Using BISELLO 88b and MANE 82 data.  
<sup>7</sup> From global fit including  $\rho$ ,  $\omega$ ,  $\phi$  and  $\rho(1700)$  assume mass 1570 MeV and width 510 MeV for  $\rho$  radial excitation.  
<sup>8</sup> From global fit of  $\rho$ ,  $\omega$ ,  $\phi$  and their radial excitations to channels  $\omega\pi^+\pi^-$ ,  $K^+K^-$ ,  $K_S^0\bar{K}_L^0$ ,  $K_S^0K^\pm\pi^\mp$ . Assume mass 1570 MeV and width 510 MeV for  $\rho$  radial excitations, mass 1570 and width 500 MeV for  $\omega$  radial excitation.  
<sup>9</sup> Fit to one channel only, neglecting interference with  $\omega$ ,  $\rho(1700)$ .

**PHOTOPRODUCTION**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1753 ± 3	<sup>10</sup> LINK	02k	FOCS 20-160 $\gamma p \rightarrow K^+K^-p$
1726 ± 22	<sup>10</sup> BUSENITZ	89	TPS $\gamma p \rightarrow K^+K^-X$
1760 ± 20	<sup>10</sup> ATKINSON	85c	OMEG 20-70 $\gamma p \rightarrow K\bar{K}X$
1690 ± 10	<sup>10</sup> ASTON	81f	OMEG 25-70 $\gamma p \rightarrow K^+K^-X$
<sup>10</sup> We list here a state decaying into $K^+K^-$ possibly different from $\phi(1680)$ .			

 **$p\bar{p}$  ANNIHILATION**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1700 ± 8	<sup>11</sup> AMSLER	06	CBAR 0.9 $\bar{p}p \rightarrow K^+K^-\pi^0$
<sup>11</sup> Could also be $\rho(1700)$ .			

 **$\phi(1680)$  WIDTH** **$e^+e^-$  PRODUCTION**

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>150±50 OUR ESTIMATE</b> This is only an educated guess; the error given is larger than the error on the average of the published values.				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
211 ± 14 ± 19	4.8k	<sup>12</sup> SHEN	09	BELL 10.6 $e^+e^- \rightarrow K^+K^-\pi^+\pi^-\gamma$
322 ± 77 ± 160		<sup>13</sup> AUBERT	08s	BABR 10.6 $e^+e^- \rightarrow \text{hadrons}$
139 ± 60	948	<sup>14</sup> AKHMETSHIN 03	CMD2	1.05-1.38 $e^+e^- \rightarrow K_L^0\bar{K}_S^0$
300 ± 60		<sup>15</sup> CLEGG	94	RVUE $e^+e^- \rightarrow K^+K^-, K_S^0\bar{K}_S^0K\pi$
146 ± 55	367	BISELLO	91c	DM2 $e^+e^- \rightarrow K_S^0K^\pm\pi^\mp$
207 ± 45		<sup>16</sup> BISELLO	88b	DM2 $e^+e^- \rightarrow K^+K^-$
185 ± 22		<sup>17</sup> BUON	82	DM1 $e^+e^- \rightarrow \text{hadrons}$
102 ± 36		<sup>18</sup> MANE	82	DM1 $e^+e^- \rightarrow K_S^0K\pi$
<sup>12</sup> From a fit with two incoherent Breit-Wigners.				
<sup>13</sup> From the simultaneous fit to the $K\bar{K}^*(892) + \text{c.c.}$ and $\phi\eta$ data from AUBERT 08s using the results of AUBERT 07AK.				
<sup>14</sup> From the combined fit of AKHMETSHIN 03 and MANE 81 also including $\rho$ , $\omega$ , and $\phi$ . Neither isospin nor flavor structure known.				
<sup>15</sup> Using BISELLO 88b and MANE 82 data.				
<sup>16</sup> From global fit including $\rho$ , $\omega$ , $\phi$ and $\rho(1700)$				
<sup>17</sup> From global fit of $\rho$ , $\omega$ , $\phi$ and their radial excitations to channels $\omega\pi^+\pi^-$ , $K^+K^-$ , $K_S^0\bar{K}_L^0$ , $K_S^0K^\pm\pi^\mp$ . Assume mass 1570 MeV and width 510 MeV for $\rho$ radial excitations, mass 1570 and width 500 MeV for $\omega$ radial excitation.				
<sup>18</sup> Fit to one channel only, neglecting interference with $\omega$ , $\rho(1700)$ .				

**PHOTOPRODUCTION**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
122 ± 63	<sup>19</sup> LINK	02k	FOCS 20-160 $\gamma p \rightarrow K^+K^-p$
121 ± 47	<sup>19</sup> BUSENITZ	89	TPS $\gamma p \rightarrow K^+K^-X$
80 ± 40	<sup>19</sup> ATKINSON	85c	OMEG 20-70 $\gamma p \rightarrow K\bar{K}X$
100 ± 40	<sup>19</sup> ASTON	81f	OMEG 25-70 $\gamma p \rightarrow K^+K^-X$
<sup>19</sup> We list here a state decaying into $K^+K^-$ possibly different from $\phi(1680)$ .			

 **$p\bar{p}$  ANNIHILATION**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
143 ± 24	<sup>20</sup> AMSLER	06	CBAR 0.9 $\bar{p}p \rightarrow K^+K^-\pi^0$
<sup>20</sup> Could also be $\rho(1700)$ .			

 **$\phi(1680)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\bar{K}^*(892) + \text{c.c.}$	dominant
$\Gamma_2$ $K_S^0\bar{K}_L^0K\pi$	seen
$\Gamma_3$ $K\bar{K}$	seen
$\Gamma_4$ $K_L^0\bar{K}_S^0$	
$\Gamma_5$ $e^+e^-$	seen

$\Gamma_6$ $\omega\pi\pi$	not seen
$\Gamma_7$ $\phi\pi\pi$	
$\Gamma_8$ $K^+K^-\pi^+\pi^-$	seen
$\Gamma_9$ $\eta\phi$	seen
$\Gamma_{10}$ $\eta\gamma$	seen
$\Gamma_{11}$ $K^+K^-\pi^0$	

 **$\phi(1680)$   $\Gamma(i)\Gamma(e^+e^-)/\Gamma^2(\text{total})$** 

This combination of a branching ratio into channel (*i*) and branching ratio into  $e^+e^-$  is directly measured and obtained from the cross section at the peak. We list only data that have not been used to determine the branching ratio into (*i*) or  $e^+e^-$ .

 **$\Gamma(K_L^0\bar{K}_S^0)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**  **$\Gamma_4/\Gamma \times \Gamma_5/\Gamma$** 

VALUE (units $10^{-6}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.131 ± 0.059	948	<sup>21</sup> AKHMETSHIN 03	CMD2	1.05-1.38 $e^+e^- \rightarrow K_L^0\bar{K}_S^0$
<sup>21</sup> From the combined fit of AKHMETSHIN 03 and MANE 81 also including $\rho$ , $\omega$ , and $\phi$ . Neither isospin nor flavor structure known. Recalculated by us.				

 **$\Gamma(K\bar{K}^*(892) + \text{c.c.})/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**  **$\Gamma_1/\Gamma \times \Gamma_5/\Gamma$** 

VALUE (units $10^{-6}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.15 ± 0.16 ± 0.01	22	AUBERT	08s	BABR 10.6 $e^+e^- \rightarrow K\bar{K}^*(892)\gamma + \text{c.c.}$
3.29 ± 1.57	367	<sup>23</sup> BISELLO	91c	DM2 1.35-2.40 $e^+e^- \rightarrow K_S^0K^\pm\pi^\mp$
<sup>22</sup> From the simultaneous fit to the $K\bar{K}^*(892) + \text{c.c.}$ and $\phi\eta$ data from AUBERT 08s using the results of AUBERT 07AK.				
<sup>23</sup> Recalculated by us with the published value of $B(K\bar{K}^*(892) + \text{c.c.}) \times \Gamma(e^+e^-)$ .				

 **$\Gamma(\phi\pi\pi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**  **$\Gamma_7/\Gamma \times \Gamma_5/\Gamma$** 

VALUE (units $10^{-7}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.86 ± 0.14 ± 0.21	4.8k	<sup>24</sup> SHEN	09	BELL 10.6 $e^+e^- \rightarrow K^+K^-\pi^+\pi^-\gamma$
<sup>24</sup> Multiplied by 3/2 to take into account the $\phi\pi^0\pi^0$ mode. Using $B(\phi \rightarrow K^+K^-) = (49.2 \pm 0.6)\%$ .				

 **$\Gamma(\eta\phi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$**  **$\Gamma_9/\Gamma \times \Gamma_5/\Gamma$** 

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.43 ± 0.10 ± 0.09	<sup>25</sup> AUBERT	08s	BABR 10.6 $e^+e^- \rightarrow \phi\eta\gamma$
<sup>25</sup> From the simultaneous fit to the $K\bar{K}^*(892) + \text{c.c.}$ and $\phi\eta$ data from AUBERT 08s using the results of AUBERT 07AK.			

 **$\phi(1680)$  BRANCHING RATIOS** **$\Gamma(K\bar{K}^*(892) + \text{c.c.})/\Gamma(K_S^0\bar{K}_L^0K\pi)$**  **$\Gamma_1/\Gamma_2$** 

VALUE	DOCUMENT ID	TECN	COMMENT
dominant	MANE	82	DM1 $e^+e^- \rightarrow K_S^0K^\pm\pi^\mp$

 **$\Gamma(K\bar{K})/\Gamma(K\bar{K}^*(892) + \text{c.c.})$**  **$\Gamma_3/\Gamma_1$** 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.07 ± 0.01	BUON	82	DM1 $e^+e^-$

 **$\Gamma(\omega\pi\pi)/\Gamma(K\bar{K}^*(892) + \text{c.c.})$**  **$\Gamma_6/\Gamma_1$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<0.10	BUON	82	DM1 $e^+e^-$

 **$\Gamma(\eta\phi)/\Gamma_{\text{total}}$**  **$\Gamma_9/\Gamma$** 

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
seen	35	<sup>26</sup> ACHASOV	14	SND 1.15-2.00 $e^+e^- \rightarrow \eta\gamma$
<sup>26</sup> From a phenomenological model based on vector meson dominance with $\rho(1450)$ and $\phi(1680)$ masses and widths from the PDG 12.				

 **$\Gamma(\eta\phi)/\Gamma(K\bar{K}^*(892) + \text{c.c.})$**  **$\Gamma_9/\Gamma_1$** 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
≈ 0.37	<sup>27</sup> AUBERT	08s	BABR 10.6 $e^+e^- \rightarrow \text{hadrons}$
<sup>27</sup> From the fit including data from AUBERT 07AK.			

 **$\Gamma(\eta\gamma)/\Gamma_{\text{total}}$**  **$\Gamma_{10}/\Gamma$** 

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
seen	35	<sup>28</sup> ACHASOV	14	SND 1.15-2.00 $e^+e^- \rightarrow \eta\gamma$
<sup>28</sup> From a phenomenological model based on vector meson dominance with $\rho(1450)$ and $\phi(1680)$ masses and widths from the PDG 12.				

# Meson Particle Listings

## $\phi(1680)$ , $\rho_3(1690)$

### $\phi(1680)$ REFERENCES

ACHASOV	14	PR D90 032002	M.N. Achasov <i>et al.</i>	(SND Collab.)
PDG	12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
SHEN	09	PR D80 031101	C.P. Shen <i>et al.</i>	(BELLE Collab.)
AUBERT	08S	PR D77 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AK	PR D76 012008	B. Aubert <i>et al.</i>	(BABAR Collab.)
AMSLER	06	PL B639 165	C. Amisler <i>et al.</i>	(CBAR Collab.)
AKHMETSHIN	03	PL B551 27	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
Also	PAN 65 1222	E.V. Anashkin, V.M. Aukchenko, R.R. Akhmetshin		
LINK	02K	PL B545 50	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ACHASOV	98H	PR D57 4334	N.N. Achasov, A.A. Kozhevnikov	(LANC, MCHS)
CLEGG	94	ZPHY C62 455	A.B. Clegg, A. Donnachie	(DM2 Collab.)
ANTONELLI	92	ZPHY C56 15	A. Antonelli <i>et al.</i>	(DM2 Collab.)
BISELLO	91C	ZPHY C52 227	D. Bisello <i>et al.</i>	(NOVO)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(ILL, FNAL)
BUSENITZ	89	PR D40 1	J.K. Busenitz <i>et al.</i>	(PADO, CLER, FRAS+)
BISELLO	88B	ZPHY C39 13	D. Bisello <i>et al.</i>	(NOVO)
BARKOV	87	JETPL 46 164	L.M. Barkov <i>et al.</i>	(NOVO)
ATKINSON	85C	ZPHY C27 233	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
BUON	82	PL 118B 221	J. Buon <i>et al.</i>	(LALO, MONP)
MANE	82	PL 112B 178	F. Mane <i>et al.</i>	(LALO)
ASTON	81F	PL 104B 231	D. Aston	(BONN, CERN, EPOL, GLAS, LAN C+)
IVANOV	81	PL 107B 297	P.M. Ivanov <i>et al.</i>	(NOVO)
MANE	81	PL 99B 261	F. Mane <i>et al.</i>	(ORSAY)

## $\rho_3(1690)$

$I^{G(J^{PC})} = 1^+(3^{--})$

### $\rho_3(1690)$ MASS

VALUE (MeV)	DOCUMENT ID
<b>1688.8±2.1 OUR AVERAGE</b>	Includes data from the 5 datablocks that follow this one.

### 2 $\pi$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

### 1686± 4 OUR AVERAGE

1677±14		EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow 2\pi p$
1679±11	476	BALTAY	78B	HBC	0	15 $\pi^+ p \rightarrow \pi^+ \pi^- n$
1678±12	175	<sup>1</sup> ANTIPOV	77	CIBS	0	25 $\pi^- p \rightarrow p 3\pi$
1690± 7	600	<sup>1</sup> ENGLER	74	DBC	0	6 $\pi^+ n \rightarrow \pi^+ \pi^- p$
1693± 8		<sup>2</sup> GRAYER	74	ASPK	0	17 $\pi^- p \rightarrow \pi^+ \pi^- n$
1678±12		MATTHEWS	71C	DBC	0	7 $\pi^+ N$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1734±10		<sup>3</sup> CORDEN	79	OMEG		12–15 $\pi^- p \rightarrow n 2\pi$
1692±12		<sup>2,4</sup> ESTABROOKS	75	RVUE		17 $\pi^- p \rightarrow \pi^+ \pi^- n$
1737±23		ARMENISE	70	DBC	0	9 $\pi^+ N$
1650±35	122	BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow N 2\pi$
1687±21		STUNTEBECK	70	HDBC	0	8 $\pi^- p, 5.4 \pi^+ d$
1683±13		ARMENISE	68	DBC	0	5.1 $\pi^+ d$
1670±30		GOLDBERG	65	HBC	0	6 $\pi^+ d, 8 \pi^- p$

<sup>1</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{n}$ ; see the note with the  $K^*(892)$  mass.  
<sup>2</sup> Uses same data as HYAMS 75.  
<sup>3</sup> From a phase shift solution containing a  $f_2'(1525)$  width two times larger than the  $K\bar{K}$  result.  
<sup>4</sup> From phase-shift analysis. Error takes account of spread of different phase-shift solutions.

### $K\bar{K}$ AND $K\bar{K}\pi$ MODES

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

### 1696± 4 OUR AVERAGE

1699± 5		ALPER	80	CNTR	0	62 $\pi^- p \rightarrow K^+ K^- n$
1698±12	6k	<sup>5,6</sup> MARTIN	78D	SPEC		10 $\pi p \rightarrow K_S^0 K^- p$
1692± 6		BLUM	75	ASPK	0	18.4 $\pi^- p \rightarrow n K^+ K^-$
1690±16		ADERHOLZ	69	HBC	+	8 $\pi^+ p \rightarrow K\bar{K}\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1694± 8		<sup>7</sup> COSTA...	80	OMEG		10 $\pi^- p \rightarrow K^+ K^- n$

<sup>5</sup> From a fit to  $J^P = 3^-$  partial wave.  
<sup>6</sup> Systematic error on mass scale subtracted.  
<sup>7</sup> They cannot distinguish between  $\rho_3(1690)$  and  $\omega_3(1670)$ .

### (4 $\pi$ ) $^\pm$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

### 1686± 5 OUR AVERAGE

1694± 6		<sup>8</sup> EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow p 4\pi$
1665±15	177	BALTAY	78B	HBC	+	15 $\pi^+ p \rightarrow p 4\pi$
1670±10		THOMPSON	74	HBC	+	13 $\pi^+ p$
1687±20		CASON	73	HBC	—	8,18.5 $\pi^- p$

1685±14		<sup>9</sup> CASON	73	HBC	—	8,18.5 $\pi^- p$
1680±40	144	BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow N 4\pi$
1689±20	102	<sup>9</sup> BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow N 2\rho$
1705±21		CASO	70	HBC	—	11.2 $\pi^- p \rightarrow n \rho 2\pi$

• • • We do not use the following data for averages, fits, limits, etc. • • •						
1718±10		<sup>10</sup> EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow p 4\pi$
1673± 9		<sup>11</sup> EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow p 4\pi$
1733± 9	66	<sup>9</sup> KLIGER	74	HBC	—	4.5 $\pi^- p \rightarrow p 4\pi$
1630±15		HOLMES	72	HBC	+	10–12 $K^+ p$
1720±15		BALTAY	68	HBC	+	7, 8.5 $\pi^+ p$
<sup>8</sup> From $\rho^- \rho^0$ mode, not independent of the other two EVANGELISTA 81 entries. <sup>9</sup> From $\rho^\pm \rho^0$ mode. <sup>10</sup> From $a_2(1320)^- \pi^0$ mode, not independent of the other two EVANGELISTA 81 entries. <sup>11</sup> From $a_2(1320)^0 \pi^-$ mode, not independent of the other two EVANGELISTA 81 entries.						

### $\omega\pi$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.				

### 1681± 7 OUR AVERAGE

1670±25		<sup>12</sup> ALDE	95	GAM2		38 $\pi^- p \rightarrow \omega \pi^0 n$
1690±15		EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow \omega \pi p$
1666±14		GESSAROLI	77	HBC		11 $\pi^- p \rightarrow \omega \pi p$
1686± 9		THOMPSON	74	HBC	+	13 $\pi^+ p$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1654±24		BARNHAM	70	HBC	+	10 $K^+ p \rightarrow \omega \pi X$
<sup>12</sup> Supersedes ALDE 92c.						

### $\eta\pi^+ \pi^-$ MODE

(For difficulties with MMS experiments, see the  $a_2(1320)$  mini-review in the 1973 edition.)

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.				

### 1682±12 OUR AVERAGE

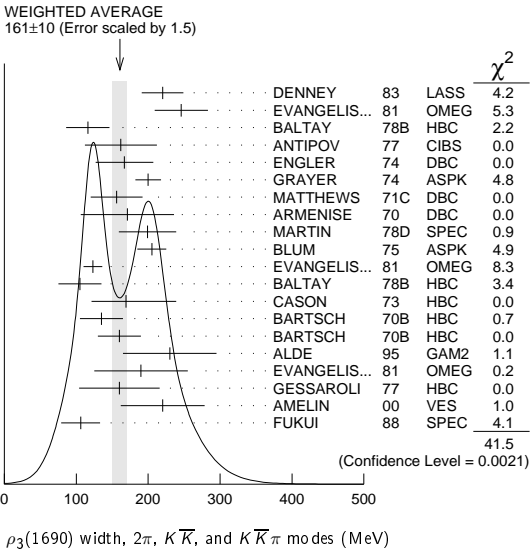
1685±10±20		AMELIN	00	VES		37 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
1680±15		FUKUI	88	SPEC	0	8.95 $\pi^- p \rightarrow \eta \pi^+ \pi^- n$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1700±47		<sup>13</sup> ANDERSON	69	MMS	—	16 $\pi^- p$ backward
1632±15		<sup>13,14</sup> FOCACCI	66	MMS	—	7–12 $\pi^- p \rightarrow p \text{MM}$
1700±15		<sup>13,14</sup> FOCACCI	66	MMS	—	7–12 $\pi^- p \rightarrow p \text{MM}$
1748±15		<sup>13,14</sup> FOCACCI	66	MMS	—	7–12 $\pi^- p \rightarrow p \text{MM}$

<sup>13</sup> Seen in 2.5–3 GeV/c  $\bar{p}p$ .  $2\pi^+ 2\pi^-$ , with 0, 1, 2  $\pi^+ \pi^-$  pairs in  $\rho$  band not seen by OREN 74 (2.3 GeV/c  $\bar{p}p$ ) with more statistics. (Jan. 1976)  
<sup>14</sup> Not seen by BOWEN 72.

### $\rho_3(1690)$ WIDTH

### 2 $\pi$ , $K\bar{K}$ , AND $K\bar{K}\pi$ MODES

VALUE (MeV)	DOCUMENT ID
<b>161±10 OUR AVERAGE</b>	Includes data from the 5 datablocks that follow this one. Error includes scale factor of 1.5. See the ideogram below.



See key on page 601

## Meson Particle Listings

 $\rho_3(1690)$ **2 $\pi$  MODE**

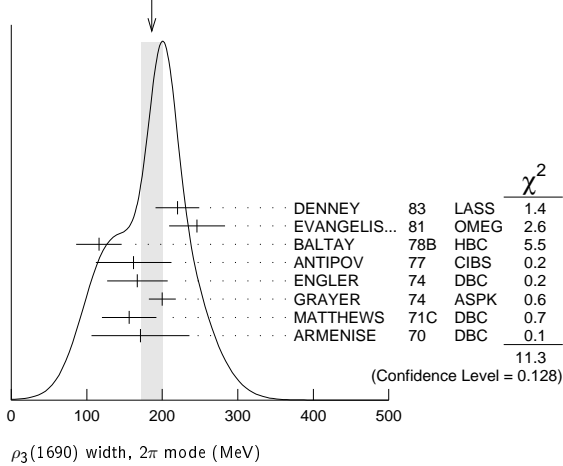
VALUE (MeV) EVTS DOCUMENT ID TECN CHG COMMENT  
The data in this block is included in the average printed for a previous datablock.

**186 $\pm$ 14 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

220 $\pm$ 29		DENNEY	83	LASS	10	$\pi^+ N$
246 $\pm$ 37		EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow 2\pi p$
116 $\pm$ 30	476	BALTAY	78B	HBC	0	15 $\pi^+ p \rightarrow$ $\pi^+ \pi^- n$
162 $\pm$ 5.0	175	15 ANTIPOV	77	CIBS	0	25 $\pi^- p \rightarrow p 3\pi$
167 $\pm$ 4.0	600	ENGLER	74	DBC	0	6 $\pi^+ n \rightarrow$ $\pi^+ \pi^- p$
200 $\pm$ 18		16 GRAYER	74	ASPK	0	17 $\pi^- p \rightarrow$ $\pi^+ \pi^- n$
156 $\pm$ 36		MATTHEWS	71C	DBC	0	7 $\pi^+ N$
171 $\pm$ 65		ARMENISE	70	DBC	0	9 $\pi^+ d$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
322 $\pm$ 35		17 CORDEN	79	OMEG	12–15	$\pi^- p \rightarrow$ $n 2\pi$
240 $\pm$ 30		16,18 ESTABROOKS	75	RVUE	17	$\pi^- p \rightarrow$ $\pi^+ \pi^- n$
180 $\pm$ 30	122	BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow N 2\pi$
267 $\pm$ 72 —46		STUNTEBECK	70	HDBC	0	8 $\pi^- p, 5.4 \pi^+ d$
188 $\pm$ 4.9		ARMENISE	68	DBC	0	5.1 $\pi^+ d$
180 $\pm$ 4.0		GOLDBERG	65	HBC	0	6 $\pi^+ d, 8 \pi^- p$

<sup>15</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.  
<sup>16</sup> Uses same data as HYAMS 75 and BECKER 79.  
<sup>17</sup> From a phase shift solution containing a  $f_2'(1525)$  width two times larger than the  $K\bar{K}$  result.  
<sup>18</sup> From phase-shift analysis. Error takes account of spread of different phase-shift solutions.

WEIGHTED AVERAGE  
186 $\pm$ 14 (Error scaled by 1.3)

 **$K\bar{K}$  AND  $K\bar{K}\pi$  MODES**

VALUE (MeV) EVTS DOCUMENT ID TECN CHG COMMENT  
The data in this block is included in the average printed for a previous datablock.

**204 $\pm$ 18 OUR AVERAGE**

199 $\pm$ 4.0	6000	19 MARTIN	78D	SPEC	10	$\pi p \rightarrow$ $K_S^0 K^- p$
205 $\pm$ 20		BLUM	75	ASPK	0	18.4 $\pi^- p \rightarrow$ $n K^+ K^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
219 $\pm$ 4		ALPER	80	CNTR	0	62 $\pi^- p \rightarrow$ $K^+ K^- n$
186 $\pm$ 11		20 COSTA...	80	OMEG	10	$\pi^- p \rightarrow$ $K^+ K^- n$
112 $\pm$ 60		ADERHOLZ	69	HBC	+	8 $\pi^+ p \rightarrow K\bar{K}\pi$

<sup>19</sup> From a fit to  $J^P = 3^-$  partial wave.  
<sup>20</sup> They cannot distinguish between  $\rho_3(1690)$  and  $\omega_3(1670)$ .

**(4 $\pi$ ) $^\pm$  MODE**

VALUE (MeV) EVTS DOCUMENT ID TECN CHG COMMENT  
The data in this block is included in the average printed for a previous datablock.

**129 $\pm$ 10 OUR AVERAGE**

123 $\pm$ 13		21 EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow p 4\pi$
105 $\pm$ 30	177	BALTAY	78B	HBC	+	15 $\pi^+ p \rightarrow p 4\pi$
169 $\pm$ 70 —48		CASON	73	HBC	—	8,18.5 $\pi^- p$
135 $\pm$ 30	144	BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow N 4\pi$
160 $\pm$ 30	102	BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow N 2\rho$

• • • We do not use the following data for averages, fits, limits, etc. • • •

230 $\pm$ 28		22 EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow p 4\pi$
184 $\pm$ 33		23 EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow p 4\pi$
150	66	24 KLIGER	74	HBC	—	4.5 $\pi^- p \rightarrow$ $p 4\pi$
106 $\pm$ 25		THOMPSON	74	HBC	+	13 $\pi^+ p$
125 $\pm$ 83 —35		24 CASON	73	HBC	—	8,18.5 $\pi^- p$
130 $\pm$ 30		HOLMES	72	HBC	+	10–12 $K^+ p$
180 $\pm$ 30	90	24 BARTSCH	70B	HBC	+	8 $\pi^+ p \rightarrow$ $N a_2 \pi$
100 $\pm$ 35		BALTAY	68	HBC	+	7, 8.5 $\pi^+ p$

<sup>21</sup> From  $\rho^- \rho^0$  mode, not independent of the other two EVANGELISTA 81 entries.

<sup>22</sup> From  $a_2(1320)^0 \pi^0$  mode, not independent of the other two EVANGELISTA 81 entries.

<sup>23</sup> From  $a_2(1320)^0 \pi^-$  mode, not independent of the other two EVANGELISTA 81 entries.

<sup>24</sup> From  $\rho^\pm \rho^0$  mode.

 **$\omega\pi$  MODE**

VALUE (MeV) DOCUMENT ID TECN CHG COMMENT  
The data in this block is included in the average printed for a previous datablock.

**190 $\pm$ 40 OUR AVERAGE**

230 $\pm$ 65		25 ALDE	95	GAM2	38	$\pi^- p \rightarrow$ $\omega \pi^0 n$
190 $\pm$ 65		EVANGELIS...	81	OMEG	—	12 $\pi^- p \rightarrow \omega \pi p$
160 $\pm$ 5.6		GESSAROLI	77	HBC		11 $\pi^- p \rightarrow \omega \pi p$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
89 $\pm$ 25		THOMPSON	74	HBC	+	13 $\pi^+ p$
130 $\pm$ 73 —43		BARNHAM	70	HBC	+	10 $K^+ p \rightarrow$ $\omega \pi X$

<sup>25</sup> Supersedes ALDE 92c.

 **$\eta\pi^+\pi^-$  MODE**

(For difficulties with MMS experiments, see the  $a_2(1320)$  mini-review in the 1973 edition.)

VALUE (MeV) DOCUMENT ID TECN CHG COMMENT  
The data in this block is included in the average printed for a previous datablock.

**126 $\pm$ 40 OUR AVERAGE** Error includes scale factor of 1.8.

220 $\pm$ 30 $\pm$ 50		AMELIN	00	VES	37	$\pi^- p \rightarrow$ $\eta \pi^+ \pi^- n$
106 $\pm$ 27		FUKUI	88	SPEC	0	8.95 $\pi^- p \rightarrow$ $\eta \pi^+ \pi^- n$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
195		26 ANDERSON	69	MMS	—	16 $\pi^- p$ back- ward
< 21	26,27	FOCACCI	66	MMS	—	7–12 $\pi^- p \rightarrow$ $p MM$
< 30	26,27	FOCACCI	66	MMS	—	7–12 $\pi^- p \rightarrow$ $p MM$
< 38	26,27	FOCACCI	66	MMS	—	7–12 $\pi^- p \rightarrow$ $p MM$

<sup>26</sup> Seen in 2.5–3 GeV/c  $\bar{p}p$ .  $2\pi^+ 2\pi^-$ , with 0, 1, 2  $\pi^+ \pi^-$  pairs in  $\rho^0$  band not seen by OREN 74 (2.3 GeV/c  $\bar{p}p$ ) with more statistics. (Jan. 1979)

<sup>27</sup> Not seen by BOWEN 72.

 **$\rho_3(1690)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor
$\Gamma_1$ 4 $\pi$	(71.1 $\pm$ 1.9 ) %	
$\Gamma_2$ $\pi^\pm \pi^+ \pi^- \pi^0$	(67 $\pm$ 22 ) %	
$\Gamma_3$ $\omega \pi$	(16 $\pm$ 6 ) %	
$\Gamma_4$ $\pi \pi$	(23.6 $\pm$ 1.3 ) %	
$\Gamma_5$ $K\bar{K}\pi$	( 3.8 $\pm$ 1.2 ) %	
$\Gamma_6$ $K\bar{K}$	( 1.58 $\pm$ 0.26 ) %	1.2
$\Gamma_7$ $\eta \pi^+ \pi^-$	seen	
$\Gamma_8$ $\rho(770)\eta$	seen	
$\Gamma_9$ $\pi \pi \rho$	seen	
Excluding $2\rho$ and $a_2(1320)\pi$ .		
$\Gamma_{10}$ $a_2(1320)\pi$	seen	
$\Gamma_{11}$ $\rho \rho$	seen	
$\Gamma_{12}$ $\phi \pi$		
$\Gamma_{13}$ $\eta \pi$		
$\Gamma_{14}$ $\pi^\pm 2\pi^+ 2\pi^- \pi^0$		

# Meson Particle Listings

## $\rho_3(1690)$

### CONSTRAINED FIT INFORMATION

An overall fit to 5 branching ratios uses 10 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 14.7$  for 7 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_4$	−77		
$x_5$	−74	17	
$x_6$	−15	2	0
	$x_1$	$x_4$	$x_5$

### $\rho_3(1690)$ BRANCHING RATIOS

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_4/\Gamma$
<b>0.236±0.013 OUR FIT</b>					
<b>0.243±0.013 OUR AVERAGE</b>					
0.259 <sup>+0.018</sup> <sub>−0.019</sub>	BECKER	79	ASPK	0	17 $\pi^- p$ polarized
0.23 ± 0.02	CORDEN	79	OMEG		12–15 $\pi^- p \rightarrow n 2\pi$
0.22 ± 0.04	<sup>28</sup> MATTHEWS	71c	HDHC	0	7 $\pi^+ n \rightarrow \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.245 ± 0.006	<sup>29</sup> ESTABROOKS	75	RVUE		17 $\pi^- p \rightarrow \pi^+ \pi^- n$

<sup>28</sup> One-pion-exchange model used in this estimation.

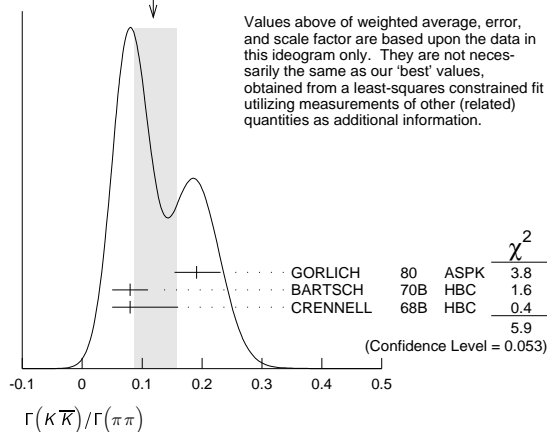
<sup>29</sup> From phase-shift analysis of HYAMS 75 data.

$\Gamma(\pi\pi)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_4/\Gamma_2$
<b>0.35±0.11</b>	CASON	73	HBC	−	8,18.5 $\pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.2	HOLMES	72	HBC	+	10–12 $K^+ p$
<0.12	BALLAM	71b	HBC	−	16 $\pi^- p$

$\Gamma(\pi\pi)/\Gamma(4\pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_4/\Gamma_1$
<b>0.332±0.026 OUR FIT</b>					
<b>0.30 ± 0.10</b>	BALTAY	78b	HBC	0	15 $\pi^+ p \rightarrow p 4\pi$

$\Gamma(K\bar{K})/\Gamma(\pi\pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_6/\Gamma_4$
<b>0.067±0.011 OUR FIT</b>					
<b>0.118<sup>+0.040</sup><sub>−0.032</sub> OUR AVERAGE</b>					
below.					
0.191 <sup>+0.040</sup> <sub>−0.037</sub>	GORLICH	80	ASPK	0	17,18 $\pi^- p$ polarized
0.08 ± 0.03	BARTSCH	70b	HBC	+	8 $\pi^+ p$
0.08 <sup>+0.08</sup> <sub>−0.03</sub>	CRENNELL	68b	HBC		6.0 $\pi^- p$

WEIGHTED AVERAGE  
0.118±0.040-0.032 (Error scaled by 1.7)



$\Gamma(K\bar{K}\pi)/\Gamma(\pi\pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_5/\Gamma_4$
<b>0.16±0.05 OUR FIT</b>					
<b>0.16±0.05</b>	<sup>30</sup> BARTSCH	70b	HBC	+	8 $\pi^+ p$

<sup>30</sup> Increased by us to correspond to  $B(\rho_3(1690) \rightarrow \pi\pi) = 0.24$ .

$[\Gamma(\pi\pi\rho) + \Gamma(a_2(1320)\pi) + \Gamma(\rho\rho)]/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$(\Gamma_9+\Gamma_{10}+\Gamma_{11})/\Gamma_2$
<b>0.94±0.09 OUR AVERAGE</b>					
0.96±0.21	BALTAY	78b	HBC	+	15 $\pi^+ p \rightarrow p 4\pi$
0.88±0.15	BALLAM	71b	HBC	−	16 $\pi^- p$
1 ± 0.15	BARTSCH	70b	HBC	+	8 $\pi^+ p$
consistent with 1	CASO	68	HBC	−	11 $\pi^- p$

$\Gamma(\rho\rho)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{11}/\Gamma_2$
<b>0.12±0.11</b>	BALTAY	78b	HBC	+	15 $\pi^+ p \rightarrow p 4\pi$
0.56	KLIGER	74	HBC	−	4.5 $\pi^- p \rightarrow p 4\pi$
0.13±0.09	<sup>31</sup> THOMPSON	74	HBC	+	13 $\pi^+ p$
0.7 ± 0.15	BARTSCH	70b	HBC	+	8 $\pi^+ p$
<sup>31</sup> $\rho\rho$ and $a_2(1320)\pi$ modes are indistinguishable.					

$\Gamma(\rho\rho)/[\Gamma(\pi\pi\rho) + \Gamma(a_2(1320)\pi) + \Gamma(\rho\rho)]$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{11}/(\Gamma_9+\Gamma_{10}+\Gamma_{11})$
<b>0.48±0.16</b>	CASO	68	HBC	−	11 $\pi^- p$

$\Gamma(a_2(1320)\pi)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{10}/\Gamma_2$
<b>0.66±0.08</b>	BALTAY	78b	HBC	+	15 $\pi^+ p \rightarrow p 4\pi$
0.36±0.14	<sup>32</sup> THOMPSON	74	HBC	+	13 $\pi^+ p$
not seen	CASON	73	HBC	−	8,18.5 $\pi^- p$
0.6 ± 0.15	BARTSCH	70b	HBC	+	8 $\pi^+ p$
0.6	BALTAY	68	HBC	+	7,8.5 $\pi^+ p$
<sup>32</sup> $\rho\rho$ and $a_2(1320)\pi$ modes are indistinguishable.					

$\Gamma(\omega\pi)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_3/\Gamma_2$
<b>0.23±0.05 OUR AVERAGE</b>					
0.33±0.07	THOMPSON	74	HBC	+	13 $\pi^+ p$
0.12±0.07	BALLAM	71b	HBC	−	16 $\pi^- p$
0.25±0.10	BALTAY	68	HBC	+	7,8.5 $\pi^+ p$
0.25±0.10	JOHNSTON	68	HBC	−	7.0 $\pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.11	BALTAY	78b	HBC	+	15 $\pi^+ p \rightarrow p 4\pi$
<0.09	KLIGER	74	HBC	−	4.5 $\pi^- p \rightarrow p 4\pi$

$\Gamma(\phi\pi)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{12}/\Gamma_2$
<b>&lt;0.11</b>	BALTAY	68	HBC	+	7,8.5 $\pi^+ p$

$\Gamma(\pi^\pm 2\pi^+ 2\pi^- \pi^0)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{14}/\Gamma_2$
<b>&lt;0.15</b>	BALTAY	68	HBC	+	7,8.5 $\pi^+ p$

$\Gamma(\eta\pi)/\Gamma(\pi^\pm\pi^\pm\pi^-\pi^0)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_{13}/\Gamma_2$
<b>&lt;0.02</b>	THOMPSON	74	HBC	+	13 $\pi^+ p$

$\Gamma(K\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_6/\Gamma$
<b>0.0158±0.0026 OUR FIT</b>					
<b>0.0130±0.0024 OUR AVERAGE</b>					
0.013 ± 0.003	COSTA...	80	OMEG	0	10 $\pi^- p \rightarrow K^+ K^- n$
0.013 ± 0.004	<sup>33</sup> MARTIN	78b	SPEC	−	10 $\pi p \rightarrow K_S^0 K^- p$

<sup>33</sup> From  $(\Gamma_4\Gamma_6)^{1/2} = 0.056 \pm 0.034$  assuming  $B(\rho_3(1690) \rightarrow \pi\pi) = 0.24$ .

$\Gamma(\omega\pi)/[\Gamma(\omega\pi) + \Gamma(\rho\rho)]$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_3/(\Gamma_3+\Gamma_{11})$
<b>0.22±0.08</b>	CASON	73	HBC	−	8,18.5 $\pi^- p$

$\Gamma(\eta\pi^+\pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
<b>seen</b>	FUKUI	88	SPEC	8.95 $\pi^- p \rightarrow \eta\pi^+\pi^- n$



See key on page 601

## Meson Particle Listings

 $\rho_3(1690)$ ,  $\rho(1700)$ 

$\Gamma(\rho_3(1320)\pi)/\Gamma(\rho(770)\eta)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma_8$
VALUE				
$5.5 \pm 2.0$	AMELIN	00	VES	$37 \pi^- p \rightarrow \eta \pi^+ \pi^- n$

 $\rho_3(1690)$  REFERENCES

AMELIN	00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
ALDE	95	ZPHY C66 379	D.M. Alde <i>et al.</i>	(GAMS Collab.) JP
ALDE	92C	ZPHY C54 553	D.M. Alde <i>et al.</i>	(BELG, SERP, KEK, LANL+)
FUKUI	88	PL B202 441	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
DENNEY	83	PR D28 2726	D.L. Denney <i>et al.</i>	(IOWA, MICH)
EVANGELIS...	81	NP B178 197	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
ALPER	80	PL 94B 422	B. Alper <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
COSTA...	80	NP B175 402	G. Costa de Beauregard <i>et al.</i>	(BARI, BONN+)
GÖRlich	80	NP B174 16	L. Görlich <i>et al.</i>	(CRAC, MPIM, CERN+)
BECKER	79	NP B151 46	H. Becker <i>et al.</i>	(MPIM, CERN, ZEEM, CRAC)
CORDEN	79	NP B157 250	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TELA+)
BALTAY	78B	PR D17 62	C. Baltay <i>et al.</i>	(COLU, BIRG)
MARTIN	78B	NP B140 158	A.D. Martin <i>et al.</i>	(DURH, GEVA)
MARTIN	78D	PL 74B 417	A.D. Martin <i>et al.</i>	(DURH, GEVA)
ANTIPOV	77	NP B119 45	Y.M. Antipov <i>et al.</i>	(SERP, GEVA)
GESSAROLI	77	NP B126 382	R. Gessaroli <i>et al.</i>	(BGNA, FIRZ, GENO+)
BLUM	75	PL 57B 403	W. Blum <i>et al.</i>	(CERN, MPIM) JP
ESTABROOKS	75	NP B95 322	P.G. Estabrooks, A.D. Martin	(DURH)
HYAMS	75	NP B100 205	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
ENGLER	74	PR D10 2070	A. Engler <i>et al.</i>	(CMU, CASE)
GRAYER	74	NP B75 189	G. Grayer <i>et al.</i>	(CERN, MPIM)
KLIGER	74	SJNP 19 428	G.K. Kliger <i>et al.</i>	(ITEP)
		Translated from YAF 19 839		
OREN	74	NP B71 189	Y. Oren <i>et al.</i>	(ANL, OXF)
THOMPSON	74	NP B69 220	G. Thompson <i>et al.</i>	(PURD)
CASON	73	PR D7 1971	N.M. Cason <i>et al.</i>	(NDAM)
BOWEN	72	PRL 29 890	D.R. Bowen <i>et al.</i>	(NEAS, STON)
HOLMES	72	PR D6 3336	R. Holmes <i>et al.</i>	(ROCH)
BALLAM	71B	PR D3 2606	J. Ballam <i>et al.</i>	(SLAC)
MATTHEWS	71C	NP B33 1	J.A.J. Matthews <i>et al.</i>	(TNT0, WISC) JP
ARMENISE	70	LNC 4 199	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ)
BARNHAM	70	PRL 24 1083	K.W.J. Barnham <i>et al.</i>	(BIRM)
BARTSCH	70B	NP B22 109	J. Bartsch <i>et al.</i>	(AACH, BERL, CERN)
CASO	70	LNC 3 707	C. Caso <i>et al.</i>	(GENO, HAMB, MILA, SACL)
STUNTEBECK	70	PL 32B 391	PH. Stuntebeck <i>et al.</i>	(NDAM)
ADERHOLZ	69	NP B11 259	M. Aderholz <i>et al.</i>	(AACH3, BERL, CERN+)
ANDERSON	69	PRL 22 1390	E.W. Anderson <i>et al.</i>	(BNL, CMU)
ARMENISE	68	NC 54A 999	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ+)
BALTAY	68	PRL 20 887	C. Baltay <i>et al.</i>	(COLU, ROCH, RUTG, YALE)
CASO	68	NC 54A 983	C. Caso <i>et al.</i>	(GENO, HAMB, MILA, SACL)
CRENNELL	68B	PL 28B 136	D.J. Crennell <i>et al.</i>	(BNL)
JOHNSTON	68	PRL 20 1414	T.F. Johnston <i>et al.</i>	(TNT0, WISC) JP
FOCACCI	66	PRL 17 890	M.N. Focacci <i>et al.</i>	(CERN)
GOLDBERG	65	PL 17 354	M. Goldberg <i>et al.</i>	(CERN, EPOL, ORSAY+)

 $\rho(1700)$ 

$$J^G(J^{PC}) = 1^+(1^{--})$$

THE  $\rho(1450)$  AND THE  $\rho(1700)$ 

Updated November 2015 by S. Eidelman (Novosibirsk), C. Hanhart (Juelich) and G. Venanzoni (Frascati).

In our 1988 edition, we replaced the  $\rho(1600)$  entry with two new ones, the  $\rho(1450)$  and the  $\rho(1700)$ , because there was emerging evidence that the 1600-MeV region actually contains two  $\rho$ -like resonances. Erkal [1] had pointed out this possibility with a theoretical analysis on the consistency of  $2\pi$  and  $4\pi$  electromagnetic form factors and the  $\pi\pi$  scattering length. Donnachie [2], with a full analysis of data on the  $2\pi$  and  $4\pi$  final states in  $e^+e^-$  annihilation and photoproduction reactions, had also argued that in order to obtain a consistent picture, two resonances were necessary. The existence of  $\rho(1450)$  was supported by the analysis of  $\eta\rho^0$  mass spectra obtained in photoproduction and  $e^+e^-$  annihilation [3], as well as that of  $e^+e^- \rightarrow \omega\pi$  [4].

The analysis of [2] was further extended by [5,6] to include new data on  $4\pi$ -systems produced in  $e^+e^-$  annihilation, and in  $\tau$ -decays ( $\tau$  decays to  $4\pi$ , and  $e^+e^-$  annihilation to  $4\pi$  can be related by the Conserved Vector Current assumption). These systems were successfully analyzed using interfering contributions from two  $\rho$ -like states, and from the tail of the  $\rho(770)$  decaying into two-body states. While specific conclusions on  $\rho(1450) \rightarrow 4\pi$  were obtained, little could be said about the  $\rho(1700)$ .

Independent evidence for two  $1^-$  states is provided by [7] in  $4\pi$  electroproduction at  $\langle Q^2 \rangle = 1$  (GeV/c)<sup>2</sup>, and by [8] in a high-statistics sample of the  $\eta\pi\pi$  system in  $\pi^-p$  charge exchange.

This scenario with two overlapping resonances is supported by other data. Bisello [9] measured the pion form factor in the interval 1.35–2.4 GeV, and observed a deep minimum around 1.6 GeV. The best fit was obtained with the hypothesis of  $\rho$ -like resonances at 1420 and 1770 MeV, with widths of about 250 MeV. Antonelli [10] found that the  $e^+e^- \rightarrow \eta\pi^+\pi^-$  cross section is better fitted with two fully interfering Breit-Wigners, with parameters in fair agreement with those of [2] and [9]. These results can be considered as a confirmation of the  $\rho(1450)$ .

Decisive evidence for the  $\pi\pi$  decay mode of both  $\rho(1450)$  and  $\rho(1700)$  comes from  $\bar{p}p$  annihilation at rest [11]. It has been shown that these resonances also possess a  $K\bar{K}$  decay mode [12–14]. High-statistics studies of the decays  $\tau \rightarrow \pi\pi\nu_\tau$  [15,16], and  $\tau \rightarrow 4\pi\nu_\tau$  [17] also require the  $\rho(1450)$ , but are not sensitive to the  $\rho(1700)$ , because it is too close to the  $\tau$  mass. A recent very-high-statistics study of the  $\tau \rightarrow \pi\pi\nu_\tau$  decay performed at Belle [18] reports the first observation of both  $\rho(1450)$  and  $\rho(1700)$  in  $\tau$  decays. A clear picture of the two  $\pi^+\pi^-$  resonances interfering with the  $\rho(770)$  was also reported by BaBar using the ISR method [19].

The structure of these  $\rho$  states is not yet completely clear. Barnes [20] and Close [21] claim that  $\rho(1450)$  has a mass consistent with radial  $2S$ , but its decays show characteristics of hybrids, and suggest that this state may be a  $2S$ -hybrid mixture. Donnachie [22] argues that hybrid states could have a  $4\pi$  decay mode dominated by the  $a_1\pi$ . Such behavior has been observed by [23] in  $e^+e^- \rightarrow 4\pi$  in the energy range 1.05–1.38 GeV, and by [17] in  $\tau \rightarrow 4\pi$  decays. CLEO [24] and Belle [25] observe the  $\rho(1450) \rightarrow \omega\pi$  decay mode in  $B$ -meson decays, however, do not find  $\rho(1700) \rightarrow \omega\pi^0$ . A similar conclusion is made by [26], who studied the process  $e^+e^- \rightarrow \omega\pi^0$ . Various decay modes of the  $\rho(1450)$  and  $\rho(1700)$  are observed in  $\bar{p}n$  and  $\bar{p}p$  annihilation [27,28], but no definite conclusions can be drawn. More data should be collected to clarify the nature of the  $\rho$  states, particularly in the energy range above 1.6 GeV.

We now list under a separate entry the  $\rho(1570)$ , the  $\phi\pi$  state with  $J^{PC} = 1^{--}$  earlier observed by [29] (referred to as  $C(1480)$ ) and recently confirmed by [30]. While [31] shows that it may be a threshold effect, [5] and [32] suggest two independent vector states with this decay mode. The  $C(1480)$  has not been seen in the  $\bar{p}p$  [33] and  $e^+e^-$  [34,35] experiments. However, the sensitivity of the two latter is an order of magnitude lower than that of [30]. Note that [30] can not exclude that their observation is due to an OZI-suppressed decay mode of the  $\rho(1700)$ .

Several observations on the  $\omega\pi$  system in the 1200-MeV region [36–42] may be interpreted in terms of either  $J^P = 1^-$   $\rho(770) \rightarrow \omega\pi$  production [43], or  $J^P = 1^+$   $b_1(1235)$  production [41,42]. We argue that no special entry for a  $\rho(1250)$  is needed. The LASS amplitude analysis [44] showing

# Meson Particle Listings

## $\rho(1700)$

evidence for  $\rho(1270)$  is preliminary and needs confirmation. For completeness, the relevant observations are listed under the  $\rho(1450)$ .

Recently [45] reported a very broad  $1^{--}$  resonance-like  $K^+K^-$  state in  $J/\psi \rightarrow K^+K^-\pi^0$  decays. Its pole position corresponds to mass of 1576 MeV and width of 818 MeV. [46–48] suggest its exotic structure (molecular or multiquark), while [49] and [50] explain it by the interference between the  $\rho(1450)$  and  $\rho(1700)$ . We quote [45] as  $X(1575)$  in the section “Further States.”

Evidence for  $\rho$ -like mesons decaying into  $6\pi$  states was first noted by [51] in the analysis of  $6\pi$  mass spectra from  $e^+e^-$  annihilation [52,53] and diffractive photoproduction [54]. Clegg [51] argued that two states at about 2.1 and 1.8 GeV exist: while the former is a candidate for the  $\rho(2150)$ , the latter could be a manifestation of the  $\rho(1700)$  distorted by threshold effects. BaBar reported observations of the new decay modes of the  $\rho(2150)$  in the channels  $\eta'(958)\pi^+\pi^-$  and  $f_1(1285)\pi^+\pi^-$  [55]. The relativistic quark model [56] predicts the  $2^3D_1$  state with  $J^{PC} = 1^{--}$  at 2.15 GeV which can be identified with the  $\rho(2150)$ .

We no longer list under a separate particle  $\rho(1900)$  various observations of irregular behavior of the cross sections near the  $N\bar{N}$  threshold. Dips of various width around 1.9 GeV were reported by the E687 Collaboration (a narrow one in the  $3\pi^+3\pi^-$  diffractive photoproduction [57,58]), by the FENICE experiment (a narrow structure in the  $R$  value [59]), by BaBar in ISR (a narrow structure in  $e^+e^- \rightarrow \phi\pi$  final state [60], but much broader in  $e^+e^- \rightarrow 3\pi^+3\pi^-$  and  $e^+e^- \rightarrow 2(\pi^+\pi^-\pi^0)$  [61]), by CMD-3 (also a rather broad dip in  $e^+e^- \rightarrow 3\pi^+3\pi^-$  [62]). Most probably, these structures emerge as a threshold effect due to the opening of the  $N\bar{N}$  channel [63,64].

### References

- C. Erkal, Z. Phys. **C31**, 615 (1986).
- A. Donnachie and H. Mirzaie, Z. Phys. **C33**, 407 (1987).
- A. Donnachie and A.B. Clegg, Z. Phys. **C34**, 257 (1987).
- A. Donnachie and A.B. Clegg, Z. Phys. **C51**, 689 (1991).
- A.B. Clegg and A. Donnachie, Z. Phys. **C40**, 313 (1988).
- A.B. Clegg and A. Donnachie, Z. Phys. **C62**, 455 (1994).
- T.J. Killian *et al.*, Phys. Rev. **D21**, 3005 (1980).
- S. Fukui *et al.*, Phys. Lett. **B202**, 441 (1988).
- D. Bisello *et al.*, Phys. Lett. **B220**, 321 (1989).
- A. Antonelli *et al.*, Phys. Lett. **B212**, 133 (1988).
- A. Abele *et al.*, Phys. Lett. **B391**, 191 (1997).
- A. Abele *et al.*, Phys. Rev. **D57**, 3860 (1998).
- A. Bertin *et al.*, Phys. Lett. **B434**, 180 (1998).
- A. Abele *et al.*, Phys. Lett. **B468**, 178 (1999).
- R. Barate *et al.*, Z. Phys. **C76**, 15 (1997).
- S. Anderson, Phys. Rev. **D61**, 112002 (2000).
- K.W. Edwards *et al.*, Phys. Rev. **D61**, 072003 (2000).
- M. Fujikawa *et al.*, Phys. Rev. **D78**, 072006 (2008).
- J.P. Lees *et al.*, Phys. Rev. **D86**, 032013 (2012).
- T. Barnes *et al.*, Phys. Rev. **D55**, 4157 (1997).
- F.E. Close *et al.*, Phys. Rev. **D56**, 1584 (1997).
- A. Donnachie and Yu.S. Kalashnikova, Phys. Rev. **D60**, 114011 (1999).
- R.R. Akhmetshin *et al.*, Phys. Lett. **B466**, 392 (1999).
- J.P. Alexander *et al.*, Phys. Rev. **D64**, 092001 (2001).
- D. Matvienko *et al.*, Phys. Rev. **D92**, 012013 (2015).
- R.R. Akhmetshin *et al.*, Phys. Lett. **B562**, 173 (2003).
- A. Abele *et al.*, Eur. Phys. J. **C21**, 261 (2001).
- M. Bargiotti *et al.*, Phys. Lett. **B561**, 233 (2003).
- S.I. Bitukov *et al.*, Phys. Lett. **B188**, 383 (1987).
- B. Aubert *et al.*, Phys. Rev. **D77**, 092002 (2008).
- N.N. Achasov and G.N. Shestakov, Phys. Atom. Nucl. **59**, 1262 (1996).
- L.G. Landsberg, Sov. J. Nucl. Phys. **55**, 1051 (1992).
- A. Abele *et al.*, Phys. Lett. **B415**, 280 (1997).
- V.M. Aulchenko *et al.*, Sov. Phys. JETP Lett. **45**, 145 (1987).
- D. Bisello *et al.*, Z. Phys. **C52**, 227 (1991).
- P. Frenkiel *et al.*, Nucl. Phys. **B47**, 61 (1972).
- G. Cosme *et al.*, Phys. Lett. **B63**, 352 (1976).
- D.P. Barber *et al.*, Z. Phys. **C4**, 169 (1980).
- D. Aston, Phys. Lett. **B92**, 211 (1980).
- M. Atkinson *et al.*, Nucl. Phys. **B243**, 1 (1984).
- J.E. Brau *et al.*, Phys. Rev. **D37**, 2379 (1988).
- C. Amsler *et al.*, Phys. Lett. **B311**, 362 (1993).
- J. Layssac and F.M. Renard, Nuovo Cimento **6A**, 134 (1971).
- D. Aston *et al.*, Nucl. Phys. (Proc. Supp.) **B21**, 105 (1991).
- M. Ablikim *et al.*, Phys. Rev. Lett. **97**, 142002 (2006).
- G.-J. Ding and M.-L. Yan, Phys. Lett. **B643**, 33 (2006).
- F.K. Guo *et al.*, Nucl. Phys. **A773**, 78 (2006).
- A. Zhang *et al.*, Phys. Rev. **D76**, 036004 (2007).
- B.A. Li, Phys. Rev. **D76**, 094016 (2007).
- X. Liu *et al.*, Phys. Rev. **D75**, 074017 (2007).
- A.B. Clegg and A. Donnachie, Z. Phys. **C45**, 677 (1990).
- D. Bisello *et al.*, Phys. Lett. **107B**, 145 (1981).
- A. Castro *et al.*, LAL-88-58(1988).
- M. Atkinson *et al.*, Z. Phys. **C29**, 333 (1985).
- B. Aubert *et al.*, Phys. Rev. **D76**, 092005 (2007).
- S. Godfrey and N. Isgur, Phys. Rev. **D32**, 189 (1985).
- P.L. Frabetti *et al.*, Phys. Lett. **B514**, 240 (2001).
- P.L. Frabetti *et al.*, Phys. Lett. **B578**, 290 (2004).
- A. Antonelli *et al.*, Phys. Lett. **B365**, 427 (1996).
- B. Aubert *et al.*, Phys. Rev. **D77**, 092002 (2008).
- B. Aubert *et al.*, Phys. Rev. **D73**, 052003 (2006).
- R.R. Akhmetshin *et al.*, Phys. Lett. **B723**, 83 (2013).
- A. Obrazovsky and S. Serednyakov, JETP Lett. **99**, 315 (2014).
- J. Heidenauer *et al.*, Phys. Rev. **D92**, 054032 (2015).

### $\rho(1700)$ MASS

$\eta\rho^0$  AND  $\pi^+\pi^-$  MODES  
 VALUE (MeV)  
**1720±20 OUR ESTIMATE**

DOCUMENT ID

See key on page 601

## Meson Particle Listings

 $\rho(1700)$  $\eta\rho^0$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

1740 ± 20	ANTONELLI	88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
1701 ± 15	<sup>1</sup> FUKUI	88	SPEC	$8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

<sup>1</sup> Assuming  $\rho^+ f_0(1370)$  decay mode interferes with  $a_1(1260)^+\pi$  background. From a two Breit-Wigner fit. $\pi\pi$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

1780 ± 20	<sup>+15</sup> <sub>-20</sub>	63.5k	<sup>2</sup> ABRAMOWICZ12	ZEUS	$e p \rightarrow e\pi^+\pi^-p$
1861 ± 17			<sup>3</sup> LEES	12G	BABR $e^+e^- \rightarrow \pi^+\pi^-\gamma$
1728 ± 17	± 89	5.4M	<sup>4,5</sup> FUJIKAWA	08	BELL $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
1780	<sup>+37</sup> <sub>-29</sub>		<sup>6</sup> ABELE	97	CBAR $\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
1719 ± 15			<sup>6</sup> BERTIN	97c	OBLX $0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
1730 ± 30			CLEGG	94	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
1768 ± 21			BISELLO	89	DM2 $e^+e^- \rightarrow \pi^+\pi^-$
1745.7 ± 91.9			DUBNICKA	89	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
1546 ± 26			GESHKEN...	89	RVUE
1650			<sup>7</sup> ERKAL	85	RVUE $20-70\gamma p \rightarrow \gamma\pi$
1550 ± 70			ABE	84B	HYBR $20\gamma p \rightarrow \pi^+\pi^-p$
1590 ± 20			<sup>8</sup> ASTON	80	OMEG $20-70\gamma p \rightarrow p2\pi$
1600 ± 10			<sup>9</sup> ATIYA	79B	SPEC $50\gamma C \rightarrow C2\pi$
1598	<sup>+24</sup> <sub>-22</sub>		BECKER	79	ASPK $17\pi^-p$ polarized
1659 ± 25			<sup>7</sup> LANG	79	RVUE
1575			<sup>7</sup> MARTIN	78C	RVUE $17\pi^-p \rightarrow \pi^+\pi^-n$
1610 ± 30			<sup>7</sup> FROGGATT	77	RVUE $17\pi^-p \rightarrow \pi^+\pi^-n$
1590 ± 20			<sup>10</sup> HYAMS	73	ASPK $17\pi^-p \rightarrow \pi^+\pi^-n$

<sup>2</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.<sup>3</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.<sup>4</sup>  $|F_\pi(0)|^2$  fixed to 1.<sup>5</sup> From the GOUNARIS 68 parametrization of the pion form factor.<sup>6</sup> T-matrix pole.<sup>7</sup> From phase shift analysis of HYAMS 73 data.<sup>8</sup> Simple relativistic Breit-Wigner fit with constant width.<sup>9</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.<sup>10</sup> Included in BECKER 79 analysis. $\pi\omega$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1708 ± 41	7815	<sup>11</sup> ACHASOV	13	SND	$1.05-2.00e^+e^- \rightarrow \pi^0\pi^0\gamma$
1550 to 1620		<sup>12</sup> ACHASOV	00i	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
1580 to 1710		<sup>13</sup> ACHASOV	00i	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
1710 ± 90		ACHASOV	97	RVUE	$e^+e^- \rightarrow \omega\pi^0$

<sup>11</sup> From a phenomenological model based on vector meson dominance with the interfering  $\rho(1450)$  and  $\rho(1700)$  and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.<sup>12</sup> Taking into account both  $\rho(1450)$  and  $\rho(1700)$  contributions. Using the data of ACHASOV 00i on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .  $\rho(1450)$  mass and width fixed at 1400 MeV and 500 MeV respectively.<sup>13</sup> Taking into account the  $\rho(1700)$  contribution only. Using the data of ACHASOV 00i on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ . $K\bar{K}$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1740.8 ± 22.2	27k	<sup>14</sup> ABELE	99D	CBAR	$\pm$ $0.0\bar{p}p \rightarrow K^+K^-\pi^0$
1582 ± 36	1600	CLELAND	82B	SPEC	$\pm$ $50\pi p \rightarrow K_S^0 K^\pm p$

<sup>14</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ . $2(\pi^+\pi^-)$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1851 <sup>+27</sup> <sub>-24</sub>		ACHASOV	97	RVUE	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1570 ± 20		<sup>15</sup> CORDIER	82	DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1520 ± 30		<sup>16</sup> ASTON	81E	OMEG	$20-70\gamma p \rightarrow p4\pi$
1654 ± 25		<sup>17</sup> DIBIANCA	81	DBC	$\pi^+d \rightarrow pp2(\pi^+\pi^-)$
1666 ± 39		<sup>15</sup> BACCI	80	FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1780	34	KILLIAN	80	SPEC	$11e^-p \rightarrow 2(\pi^+\pi^-)$

1500		<sup>18</sup> ATIYA	79B	SPEC	$50\gamma C \rightarrow C4\pi^\pm$
1570 ± 60	65	<sup>19</sup> ALEXANDER	75	HBC	$7.5\gamma p \rightarrow p4\pi$
1550 ± 60		<sup>16</sup> CONVERSI	74	OSPK	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1550 ± 50	160	SCHACHT	74	STRC	$5.5-9\gamma p \rightarrow p4\pi$
1450 ± 100	340	SCHACHT	74	STRC	$9-18\gamma p \rightarrow p4\pi$
1430 ± 50	400	BINGHAM	72B	HBC	$9.3\gamma p \rightarrow p4\pi$

<sup>15</sup> Simple relativistic Breit-Wigner fit with model dependent width.<sup>16</sup> Simple relativistic Breit-Wigner fit with constant width.<sup>17</sup> One peak fit result.<sup>18</sup> Parameters roughly estimated, not from a fit.<sup>19</sup> Skew mass distribution compensated by Ross-Stodolsky factor. $\pi^+\pi^-\pi^0\pi^0$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1660 ± 30	AT KINSON	85B	OMEG	$20-70\gamma p$
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 $3(\pi^+\pi^-)$  AND  $2(\pi^+\pi^-\pi^0)$  MODES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1730 ± 34	<sup>20</sup> FRABETTI	04	E687	$\gamma p \rightarrow 3\pi^+3\pi^-p$
1783 ± 15	CLEGG	90	RVUE	$e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$

<sup>20</sup> From a fit with two resonances with the JACOB 72 continuum. $\rho(1700)$  WIDTH $\eta\rho^0$  AND  $\pi^+\pi^-$  MODES

VALUE (MeV)	DOCUMENT ID
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 $250 \pm 100$  OUR ESTIMATE $\eta\rho^0$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

150 ± 30	ANTONELLI	88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
282 ± 44	<sup>21</sup> FUKUI	88	SPEC	$8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

<sup>21</sup> Assuming  $\rho^+ f_0(1370)$  decay mode interferes with  $a_1(1260)^+\pi$  background. From a two Breit-Wigner fit. $\pi\pi$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

310 ± 30	<sup>+25</sup> <sub>-35</sub>	63.5k	<sup>22</sup> ABRAMOWICZ12	ZEUS	$e p \rightarrow e\pi^+\pi^-p$
316 ± 26			<sup>23</sup> LEES	12G	BABR $e^+e^- \rightarrow \pi^+\pi^-\gamma$
164 ± 21	<sup>+89</sup> <sub>-26</sub>	5.4M	<sup>24,25</sup> FUJIKAWA	08	BELL $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
275 ± 45			<sup>26</sup> ABELE	97	CBAR $\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
310 ± 40			<sup>26</sup> BERTIN	97c	OBLX $0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
400 ± 100			CLEGG	94	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
224 ± 22			BISELLO	89	DM2 $e^+e^- \rightarrow \pi^+\pi^-$
242.5 ± 163.0			DUBNICKA	89	RVUE $e^+e^- \rightarrow \pi^+\pi^-$
620 ± 60			GESHKEN...	89	RVUE
< 315			<sup>27</sup> ERKAL	85	RVUE $20-70\gamma p \rightarrow \gamma\pi$
280 ± 30			ABE	84B	HYBR $20\gamma p \rightarrow \pi^+\pi^-p$
230 ± 80			<sup>28</sup> ASTON	80	OMEG $20-70\gamma p \rightarrow p2\pi$
283 ± 14			<sup>29</sup> ATIYA	79B	SPEC $50\gamma C \rightarrow C2\pi$
175 ± 98			BECKER	79	ASPK $17\pi^-p$ polarized
232 ± 34			<sup>27</sup> LANG	79	RVUE
340			<sup>27</sup> MARTIN	78C	RVUE $17\pi^-p \rightarrow \pi^+\pi^-n$
300 ± 100			<sup>27</sup> FROGGATT	77	RVUE $17\pi^-p \rightarrow \pi^+\pi^-n$
180 ± 50			<sup>30</sup> HYAMS	73	ASPK $17\pi^-p \rightarrow \pi^+\pi^-n$

<sup>22</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.<sup>23</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.<sup>24</sup>  $|F_\pi(0)|^2$  fixed to 1.<sup>25</sup> From the GOUNARIS 68 parametrization of the pion form factor.<sup>26</sup> T-matrix pole.<sup>27</sup> From phase shift analysis of HYAMS 73 data.<sup>28</sup> Simple relativistic Breit-Wigner fit with constant width.<sup>29</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.<sup>30</sup> Included in BECKER 79 analysis. $K\bar{K}$  MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

187.2 ± 26.7	27k	<sup>31</sup> ABELE	99D	CBAR	$\pm$ $0.0\bar{p}p \rightarrow K^+K^-\pi^0$
265 ± 120	1600	CLELAND	82B	SPEC	$\pm$ $50\pi p \rightarrow K_S^0 K^\pm p$

<sup>31</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

# Meson Particle Listings

## $\rho(1700)$

### $2(\pi^+\pi^-)$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
510± 40		32 CORDIER	82 DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
400± 50		33 ASTON	81E OMEG	20–70 $\gamma p \rightarrow p4\pi$
400±146		34 DIBIANCA	81 DBC	$\pi^+d \rightarrow pp2(\pi^+\pi^-)$
700±160		32 BACCI	80 FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
100	34	KILLIAN	80 SPEC	11 $e^-p \rightarrow 2(\pi^+\pi^-)$
600		35 ATIYA	79B SPEC	50 $\gamma C \rightarrow C4\pi^\pm$
340±160	65	ALEXANDER	75 HBC	7.5 $\gamma p \rightarrow p4\pi$
360±100		33 CONVERSI	74 OSPK	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
400±120	160	37 SCHACHT	74 STRC	5.5–9 $\gamma p \rightarrow p4\pi$
850±200	340	37 SCHACHT	74 STRC	9–18 $\gamma p \rightarrow p4\pi$
650±100	400	BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p4\pi$
32 Simple relativistic Breit-Wigner fit with model-dependent width.				
33 Simple relativistic Breit-Wigner fit with constant width.				
34 One peak fit result.				
35 Parameters roughly estimated, not from a fit.				
36 Skew mass distribution compensated by Ross-Stodolsky factor.				
37 Width errors enlarged by us to $4\Gamma/\sqrt{N}$ ; see the note with the $K^*(892)$ mass.				

### $\pi^+\pi^-\pi^0\pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
300±50	ATKINSON	85B OMEG	20–70 $\gamma p$

### $\omega\pi^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
350 to 580	38 ACHASOV	00i SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
490 to 1040	39 ACHASOV	00i SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
38 Taking into account both $\rho(1450)$ and $\rho(1700)$ contributions. Using the data of ACHASOV 00i on $e^+e^- \rightarrow \omega\pi^0$ and of EDWARDS 00A on $\tau^- \rightarrow \omega\pi^-\nu_\tau$ . $\rho(1450)$ mass and width fixed at 1400 MeV and 500 MeV respectively.			
39 Taking into account the $\rho(1700)$ contribution only. Using the data of ACHASOV 00i on $e^+e^- \rightarrow \omega\pi^0$ and of EDWARDS 00A on $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .			

### $3(\pi^+\pi^-)$ AND $2(\pi^+\pi^-\pi^0)$ MODES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
315±100	40 FRABETTI	04 E687	$\gamma p \rightarrow 3\pi^+3\pi^-p$
285± 20	CLEGG	90 RVUE	$e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$
40 From a fit with two resonances with the JACOB 72 continuum.			

### $\rho(1700)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $4\pi$	
$\Gamma_2$ $2(\pi^+\pi^-)$	large
$\Gamma_3$ $\rho\pi\pi$	dominant
$\Gamma_4$ $\rho^0\pi^+\pi^-$	large
$\Gamma_5$ $\rho^0\pi^0\pi^0$	
$\Gamma_6$ $\rho^\pm\pi^\mp\pi^0$	large
$\Gamma_7$ $a_1(1260)\pi$	seen
$\Gamma_8$ $h_1(1170)\pi$	seen
$\Gamma_9$ $\pi(1300)\pi$	seen
$\Gamma_{10}$ $\rho\rho$	seen
$\Gamma_{11}$ $\pi^+\pi^-$	seen
$\Gamma_{12}$ $\pi\pi$	seen
$\Gamma_{13}$ $K\bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{14}$ $\eta\rho$	seen
$\Gamma_{15}$ $a_2(1320)\pi$	not seen
$\Gamma_{16}$ $K\bar{K}$	seen
$\Gamma_{17}$ $e^+e^-$	seen
$\Gamma_{18}$ $\pi^0\omega$	seen

### $\rho(1700)$ $\Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$

This combination of a partial width with the partial width into  $e^+e^-$  and with the total width is obtained from the cross-section into channel  $i$  in  $e^+e^-$  annihilation.

$\Gamma(2(\pi^+\pi^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_2\Gamma_{17}/\Gamma$		
VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.6 ± 0.2	DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
2.83±0.42	BACCI	80 FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$

### $\Gamma(\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.13	41 DIEKMANN	88 RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
0.029 $^{+0.016}_{-0.012}$	KURDADZE	83 OLYA	0.64–1.4 $e^+e^- \rightarrow \pi^+\pi^-$
41 Using total width = 220 MeV.			

### $\Gamma(K\bar{K}^*(892) + \text{c.c.}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.305±0.071	42 BIZOT	80 DM1	$e^+e^-$
42 Model dependent.			

### $\Gamma(\eta\rho) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
7±3	ANTONELLI	88 DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$

### $\Gamma(K\bar{K}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.035±0.029	43 BIZOT	80 DM1	$e^+e^-$
43 Model dependent.			

### $\Gamma(\rho\pi\pi) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.510±0.090	44 BIZOT	80 DM1	$e^+e^-$
44 Model dependent.			

### $\rho(1700)$ $\Gamma(i)/\Gamma(\text{total}) \times \Gamma(e^+e^-)/\Gamma(\text{total})$

### $\Gamma(\pi^0\omega)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.7±0.4	7815	45 ACHASOV	13 SND	1.05–2.00 $e^+e^- \rightarrow \pi^0\pi^0\gamma$
45 From a phenomenological model based on vector meson dominance with the interfering $\rho(1450)$ and $\rho(1700)$ and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.				

### $\rho(1700)$ BRANCHING RATIOS

### $\Gamma(\rho\pi\pi)/\Gamma(4\pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.28±0.06	46 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
46 $\omega\pi$ not included.			

### $\Gamma(\rho^0\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~1.0		DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
0.7 ±0.1	500	SCHACHT	74 STRC	5.5–18 $\gamma p \rightarrow p4\pi$
0.80		47 BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p4\pi$
47 The $\pi\pi$ system is in S-wave.				

### $\Gamma(\rho^0\pi^0\pi^0)/\Gamma(\rho^\pm\pi^\mp\pi^0)$

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.10	ATKINSON	85B OMEG		20–70 $\gamma p$
<0.15	ATKINSON	82 OMEG	0	20–70 $\gamma p \rightarrow p4\pi$

### $\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.16±0.05	48 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
48 $\omega\pi$ not included.			

### $\Gamma(h_1(1170)\pi)/\Gamma(4\pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.17±0.06	49 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
49 $\omega\pi$ not included.			

### $\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.30±0.10	50 ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
50 $\omega\pi$ not included.			

See key on page 601

## Meson Particle Listings

 $\rho(1700)$ ,  $a_2(1700)$  $\Gamma(\rho\rho)/\Gamma(4\pi)$   $\Gamma_{10}/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.09±0.03	<sup>51</sup> ABELE	01B	CBAR 0.0 $\bar{p}n \rightarrow 5\pi$
<sup>51</sup> $\omega\pi$ not included.			

 $\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.287 <sup>+0.043</sup> <sub>-0.042</sub>	BECKER	79	ASPK 17 $\pi^-p$ polarized
0.15 to 0.30	<sup>52</sup> MARTIN	78C	RVUE 17 $\pi^-p \rightarrow \pi^+\pi^-n$
<0.20	<sup>53</sup> COSTA...	77B	RVUE $e^+e^- \rightarrow 2\pi, 4\pi$
0.30 ±0.05	<sup>52</sup> FROGGATT	77	RVUE 17 $\pi^-p \rightarrow \pi^+\pi^-n$
<0.15	<sup>54</sup> EISENBERG	73	HBC 5 $\pi^+p \rightarrow \Delta^++2\pi$
0.25 ±0.05	<sup>55</sup> HYAMS	73	ASPK 17 $\pi^-p \rightarrow \pi^+\pi^-n$

<sup>52</sup> From phase shift analysis of HYAMS 73 data.<sup>53</sup> Estimate using unitarity, time reversal invariance, Breit-Wigner.<sup>54</sup> Estimated using one-pion-exchange model.<sup>55</sup> Included in BECKER 79 analysis. $\Gamma(\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{11}/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.13±0.05	ASTON	80	OMEG 20-70 $\gamma p \rightarrow p2\pi$
<0.14	<sup>56</sup> DAVIER	73	STRC 6-18 $\gamma p \rightarrow p4\pi$
<0.2	<sup>57</sup> BINGHAM	72B	HBC 9.3 $\gamma p \rightarrow p2\pi$
<sup>56</sup> Upper limit is estimate.			
<sup>57</sup> $2\sigma$ upper limit.			

 $\Gamma(\pi\pi)/\Gamma(4\pi)$   $\Gamma_{12}/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.16±0.04	<sup>58,59</sup> ABELE	01B	CBAR 0.0 $\bar{p}n \rightarrow 5\pi$
<sup>58</sup> Using ABELE 97.			
<sup>59</sup> $\omega\pi$ not included.			

 $\Gamma(K\bar{K}^*(892) + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
possibly seen	COAN	04	CLEO $\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau$

 $\Gamma(K\bar{K}^*(892) + \text{c.c.})/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{13}/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.15±0.03	<sup>60</sup> DELCOURT	81B	DM1 $e^+e^- \rightarrow \bar{K}K\pi$
<sup>60</sup> Assuming $\rho(1700)$ and $\omega$ radial excitations to be degenerate in mass.			

 $\Gamma(\eta\rho)/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
possibly seen		AKHMETSHIN	00D	CMD2 $e^+e^- \rightarrow \eta\pi^+\pi^-$
<0.04		DONNACHIE	87B	RVUE
<0.02	58	ATKINSON	86B	OMEG 20-70 $\gamma p$

 $\Gamma(\eta\rho)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{14}/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.123±0.027	DELCOURT	82	DM1 $e^+e^- \rightarrow \pi^+\pi^-MM$
~ 0.1	ASTON	80	OMEG 20-70 $\gamma p$

 $\Gamma(\pi^+\pi^- \text{ neutrals})/\Gamma(2(\pi^+\pi^-))$   $(\Gamma_5+\Gamma_6+0.714\Gamma_{14})/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.6±0.4	<sup>61</sup> BALLAM	74	HBC 9.3 $\gamma p$
<sup>61</sup> Upper limit. Background not subtracted.			

 $\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	AMELIN	00	VES 37 $\pi^-p \rightarrow \eta\pi^+\pi^-n$

 $\Gamma(K\bar{K})/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{16}/\Gamma_2$ 

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.015±0.010		<sup>62</sup> DELCOURT	81B	DM1	$e^+e^- \rightarrow \bar{K}K$
<0.04	95	BINGHAM	72B	HBC	0 9.3 $\gamma p$
<sup>62</sup> Assuming $\rho(1700)$ and $\omega$ radial excitations to be degenerate in mass.					

 $\Gamma(K\bar{K})/\Gamma(K\bar{K}^*(892) + \text{c.c.})$   $\Gamma_{16}/\Gamma_{13}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.052±0.026	BUON	82	DM1 $e^+e^- \rightarrow \text{hadrons}$

 $\Gamma(\pi^0\omega)/\Gamma_{\text{total}}$   $\Gamma_{18}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen		MATVIENKO	15	BELL $\bar{B}^0 \rightarrow D^{*+}\omega\pi^-$
seen	1.6k	ACHASOV	12	SND $e^+e^- \rightarrow \pi^0\pi^0\gamma$
not seen	2382	AKHMETSHIN	03B	CMD2 $e^+e^- \rightarrow \pi^0\pi^0\gamma$
seen		ACHASOV	97	RVUE $e^+e^- \rightarrow \omega\pi^0$

 $\rho(1700)$  REFERENCES

MATVIENKO	15	PR D92 012013	D. Matvienko <i>et al.</i>	(BELLE Collab.)
ACHASOV	13	PR D88 050103	M.N. Achasov <i>et al.</i>	(SND Collab.)
ABRAMOWICZ	12	EPJ C72 1869	H. Abramowicz <i>et al.</i>	(ZEUS Collab.)
ACHASOV	12	JETPL 94 734	M.N. Achasov <i>et al.</i>	
LEES	12G	PR D86 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
FUJIKAWA	08	PR D78 072006	M. Fujikawa <i>et al.</i>	(BELLE Collab.)
COAN	04	PRL 92 232001	T.E. Coan <i>et al.</i>	(CLEO Collab.)
FRABETTI	04	PL B578 290	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
AKHMETSHIN	03B	PL B562 173	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ABELE	01B	EPJ C21 261	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV	00I	PL B486 29	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	00D	PL B489 125	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AMELIN	00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
EDWARDS	00A	PR D61 072003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ABELE	99D	PL B468 178	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ABELE	97	PL B391 191	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACHASOV	97	PR D55 2663	N.N. Achasov <i>et al.</i>	(NOVM)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
CLEGG	94	ZPHY C62 455	A.B. Clegg, A. Donnachie	(LANC, MCHS)
CLEGG	90	ZPHY C45 677	A.B. Clegg, A. Donnachie	(LANC, MCHS)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
BISELLO	89	PL B220 321	D. Bisello <i>et al.</i>	(DM2 Collab.)
DUBNICKA	89	JP G15 1349	S. Dubnicka <i>et al.</i>	(JINR, SLOV)
GESHKEN	89	ZPHY C45 351	B.V. Geshkenbein	(ITEP)
ANTONELLI	88	PL B212 133	A. Antonelli <i>et al.</i>	(DM2 Collab.)
DIEKMANN	88	PRPL 159 99	B. Diekmann	(BONN)
FUKUI	88	PL B202 441	S. Fukui <i>et al.</i>	(SUGI, NAGO, KEK, KYOT+)
DONNACHIE	87B	ZPHY C34 257	A. Donnachie, A.B. Clegg	(MCHS, LANC)
ATKINSON	86B	ZPHY C30 531	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
ATKINSON	85B	ZPHY C26 499	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
ERKAL	85	ZPHY C29 485	C. Erkal, M.G. Olsson	(WISC)
ABE	84B	PRL 53 751	K. Abe <i>et al.</i>	(SLAC HFP Collab.)
KURDADZE	83	JETPL 37 733	L.M. Kurdadze <i>et al.</i>	(NOVO)
ATKINSON	82	PL 108B 55	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
BUON	82	PL 118B 221	J. Buon <i>et al.</i>	(LALO, MONP)
CLELAND	82B	NP B208 228	W.E. Cleland <i>et al.</i>	(DURH, GEVA, LAUS+)
CORDIER	82	PL 109B 129	A. Cordier <i>et al.</i>	(LALO)
DELCOURT	82	PL 113B 93	B. Delcourt <i>et al.</i>	(LALO)
ASTON	81E	NP B189 15	D. Aston	(BONN, CERN, EPOL, GLAS, LAN+)
DELCOURT	81B	Bonn Conf. 205	B. Delcourt	(ORSAY)
Albo		PL 101B 129	A. Conrier <i>et al.</i>	(LALO)
DIBIANCA	81	PR D23 595	F.A. di Bianca <i>et al.</i>	(CASE, CMU)
ASTON	80	PL 92B 215	D. Aston	(BONN, CERN, EPOL, GLAS, LAN+)
BACCI	80	PL 95B 139	C. Bacci <i>et al.</i>	(ROMA, FRAS)
BIZOT	80	Madison Conf. 546	J.C. Bizot <i>et al.</i>	(LALO, MONP)
KILLIAN	80	PR D21 3005	T.J. Killian <i>et al.</i>	(CORN)
ATIYA	79B	PRL 43 1691	M.S. Atiya <i>et al.</i>	(COLU, ILL, FNAL)
BECKER	79	NP B151 46	H. Becker <i>et al.</i>	(MPIM, CERN, ZEEM, CRAC)
LANG	79	PR D19 956	C.B. Lang, A. Mas-Parada	(GRAZ)
MARTIN	78C	ANP 114 1	A.D. Martin, M.R. Pennington	(CERN)
COSTA	77B	PL 71B 345	B. Costa de Beauregard, B. Pire, T.N. Truong	(EPOL)
FROGGATT	77	NP B129 89	C.D. Froggatt, J.L. Petersen	(GLAS, NORD)
ALEXANDER	75	PL 57B 487	G. Alexander <i>et al.</i>	(TELA)
BALLAM	74	NP B76 375	J. Ballam <i>et al.</i>	(SLAC, LBL, MPIM)
CONVERSI	74	PL 52B 493	M. Conversi <i>et al.</i>	(ROMA, FRAS)
SCHACHT	74	NP B81 205	P. Schacht <i>et al.</i>	(MPIM)
DAVIER	73	NP B58 31	M. Davier <i>et al.</i>	(SLAC)
EISENBERG	73	PL 43B 149	Y. Eisenberg <i>et al.</i>	(REHO)
HYAMS	73	NP B64 134	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
BINGHAM	72B	PL 41B 635	H.H. Bingham <i>et al.</i>	(LBL, UCB, SLAC)IGJP
JACOB	72	PR D5 1847	M. Jacob, R. Slansky	
GOUNARIS	68	PRL 21 244	G.J. Gounaris, J.J. Sakurai	

 $a_2(1700)$ 

$$I^G(J^{PC}) = 1^-(2^{++})$$

OMITTED FROM SUMMARY TABLE

 $a_2(1700)$  MASS

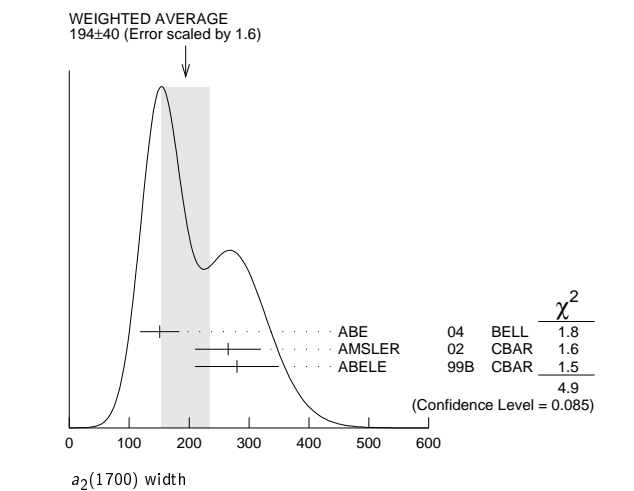
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1732±16 OUR AVERAGE</b>					Error includes scale factor of 1.9.
1737 ± 5 ± 7		ABE	04	BELL	10.6 $e^+e^- \rightarrow e^+e^-K^+K^-$
1698±44	1	AMSLER	02	CBAR	0.9 $\bar{p}p \rightarrow \pi^0\eta\eta$
1660±40		ABELE	99B	CBAR	1.94 $\bar{p}p \rightarrow \pi^0\eta\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1675±25		ANISOVICH	09	RVUE	0.0 $\bar{p}p, \pi N$
1722 ± 9 ± 15	18k	2	SCHEGELSKY	06	RVUE 0 $\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
1702 ± 7	80k	3	UMAN	06	E835 $5.2 \bar{p}p \rightarrow \eta\eta\pi^0$
1721±13±44	145k	LU	05	B852	18 $\pi^-p \rightarrow \omega\pi^-\pi^0p$
1767±14	221	4	ACCIARRI	01H	L3 $\gamma\gamma \rightarrow K_S^0 K_S^0, E_{\text{cm}}^{\text{ee}} = 91, 183-209 \text{ GeV}$
~ 1775		5	GRYGOREV	99	SPEC $40 \pi^-p \rightarrow K_S^0 K_S^0 n$
1752±21 ± 4		ACCIARRI	97T	L3	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$

Meson Particle Listings

$a_2(1700)$ ,  $f_0(1710)$

- <sup>1</sup> T-matrix pole.  
<sup>2</sup> From analysis of L3 data at 183–209 GeV.  
<sup>3</sup> Statistical error only.  
<sup>4</sup> Spin 2 dominant, isospin not determined, could also be  $I=1$ .  
<sup>5</sup> Possibly two  $J^P = 2^+$  resonances with isospins 0 and 1.

$a_2(1700)$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>194 ± 40 OUR AVERAGE</b>		Error includes scale factor of 1.6. See the ideogram below.			
151 ± 22 ± 24		ABE	04	BELL	10.6 $e^+e^- \rightarrow e^+e^-K^+K^-$
265 ± 55		<sup>6</sup> AMSLER	02	CBAR	0.9 $\overline{p}p \rightarrow \pi^0\eta\eta$
280 ± 70		ABELE	99B	CBAR	1.94 $\overline{p}p \rightarrow \pi^0\eta\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
270 <sup>+50</sup> <sub>−20</sub>		ANISOVICH	09	RVUE	0.0 $\overline{p}p, \pi N$
336 ± 20 ± 20	18k	<sup>7</sup> SCHEGELSKY	06	RVUE	0 $\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
417 ± 19	80k	<sup>8</sup> UMAN	06	E835	5.2 $\overline{p}p \rightarrow \eta\eta\pi^0$
279 ± 49 ± 66	145k	LU	05	B852	18 $\pi^-\rho \rightarrow \omega\pi^-\pi^0\rho$
187 ± 60	221	<sup>9</sup> ACCIARRI	01H	L3	$\gamma\gamma \rightarrow K_S^0 K_S^0, E_{cm}^{ee} = 91, 183\text{--}209 \text{ GeV}$
150 ± 110 ± 34		ACCIARRI	97T	L3	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
<sup>6</sup> T-matrix pole. <sup>7</sup> From analysis of L3 data at 183–209 GeV. <sup>8</sup> Statistical error only. <sup>9</sup> Spin 2 dominant, isospin not determined, could also be $I=1$ .					



$a_2(1700)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \quad \eta\pi$	seen	
$\Gamma_2 \quad \gamma\gamma$		
$\Gamma_3 \quad \rho\pi$		
$\Gamma_4 \quad f_2(1270)\pi$		
$\Gamma_5 \quad K\overline{K}$	seen	
$\Gamma_6 \quad \omega\pi^-\pi^0$	seen	
$\Gamma_7 \quad \omega\rho$	seen	

$a_2(1700)$ PARTIAL WIDTHS					
$\Gamma(\eta\pi)$					$\Gamma_1$
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
9.5 ± 2.0	870	<sup>10</sup> SCHEGELSKY	06A	RVUE	$\gamma\gamma \rightarrow K_S^0 K_S^0$
$\Gamma(\gamma\gamma)$					$\Gamma_2$
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.30 ± 0.05	870	<sup>10</sup> SCHEGELSKY	06A	RVUE	$\gamma\gamma \rightarrow K_S^0 K_S^0$
$\Gamma(K\overline{K})$					$\Gamma_5$
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
5.0 ± 3.0	870	<sup>10</sup> SCHEGELSKY	06A	RVUE	$\gamma\gamma \rightarrow K_S^0 K_S^0$
<sup>10</sup> From analysis of L3 data at 91 and 183–209 GeV, using $a_2(1700)$ mass of 1730 MeV and width of 340 MeV, and SU(3) relations.					

$a_2(1700) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$					
$[\Gamma(\rho\pi) + \Gamma(f_2(1270)\pi)] \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$			$(\Gamma_3 + \Gamma_4)\Gamma_2/\Gamma$		
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.29 ± 0.04 ± 0.02</b>		ACCIARRI	97T	L3	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.37 <sup>+0.12</sup> <sub>−0.08</sub> ± 0.10	18k	<sup>11</sup> SCHEGELSKY	06	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
$\Gamma(K\overline{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_5\Gamma_2/\Gamma$
VALUE (eV)	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
20.6 ± 4.2 ± 4.6	<sup>12</sup> ABE	04	BELL	10.6 $e^+e^- \rightarrow e^+e^-K^+K^-$	
49 ± 11 ± 13	<sup>13</sup> ACCIARRI	01H	L3	$\gamma\gamma \rightarrow K_S^0 K_S^0, E_{cm}^{ee} = 91, 183\text{--}209 \text{ GeV}$	
<sup>11</sup> From analysis of L3 data at 183–209 GeV. <sup>12</sup> Assuming spin 2. <sup>13</sup> Spin 2 dominant, isospin not determined, could also be $I=1$ .					

$a_2(1700)$ BRANCHING RATIOS					
$\Gamma(\rho\pi)/\Gamma(f_2(1270)\pi)$					$\Gamma_3/\Gamma_4$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.4 ± 0.4 ± 0.1	18k	<sup>14</sup> SCHEGELSKY	06	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
<sup>14</sup> From analysis of L3 data at 183–209 GeV.					

$a_2(1700)$ REFERENCES					
ANISOVICH	09	IJMP A24 2481	V.V. Anisovich, A.V. Sarantsev		
SCHEGELSKY	06	EPJ A27 199	V.A. Schegelsky et al.		
SCHEGELSKY	06A	EPJ A27 207	V.A. Schegelsky et al.		
UMAN	06	PR D73 052009	I. Uman et al.		(FNAL E835)
LU	05	PRL 94 032002	M. Lu et al.		(BNL E852 Collab.)
ABE	04	EPJ C32 323	K. Abe et al.		(BELLE Collab.)
AMSLER	02	EPJ C23 29	C. Amshler et al.		
ACCIARRI	01H	PL B501 173	M. Acciarri et al.		(L3 Collab.)
ABELE	99B	EPJ C8 67	A. Abele et al.		(Crystal Barrel Collab.)
GRYGOREV	99	PAN 62 470	V.K. Grygorev et al.		
		Translated from YAF 62 513.			
ACCIARRI	97T	PL B413 147	M. Acciarri et al.		(L3 Collab.)

$f_0(1710)$

$I^G(J^{PC}) = 0^+(0^{++})$

See our mini-review in the 2004 edition of this *Review*, Physics Letters **B592** 1 (2004). See also the mini-review on scalar mesons under  $f_0(500)$  (see the index for the page number).

$f_0(1710)$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1723 <sup>+6</sup> <sub>−5</sub> OUR AVERAGE</b>		Error includes scale factor of 1.6. See the ideogram below.			
1759 ± 6 <sup>+14</sup> <sub>−25</sub>	5.5k	<sup>1</sup> ABLIKIM	13N	BES3	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\eta\eta$
1750 <sup>+6</sup> <sub>−7</sub> <sup>+29</sup> <sub>−18</sub>		UEHARA	13	BELL	$\gamma\gamma \rightarrow K_S^0 K_S^0$
1701 ± 5 <sup>+9</sup> <sub>−2</sub>	4k	<sup>2</sup> CHEKANOV	08	ZEUS	$e\rho \rightarrow K_S^0 K_S^0 X$
1765 <sup>+4</sup> <sub>−3</sub> ± 13		ABLIKIM	06v	BES2	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^+\pi^-$
1760 ± 15 <sup>+15</sup> <sub>−10</sub>		<sup>3</sup> ABLIKIM	05q	BES2	$\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$
1738 ± 30		ABLIKIM	04E	BES2	$J/\psi \rightarrow \omega K^+K^-$
1740 ± 4 <sup>+10</sup> <sub>−25</sub>		<sup>4</sup> BAI	03G	BES	$J/\psi \rightarrow \gamma K\overline{K}$
1740 <sup>+30</sup> <sub>−25</sub>		<sup>4</sup> BAI	00A	BES	$J/\psi \rightarrow \gamma(\pi^+\pi^-\pi^+\pi^-)$
1698 ± 18		<sup>5</sup> BARBERIS	00E		450 $pp \rightarrow p_f\eta\eta p_S$
1710 ± 12 ± 11		<sup>6</sup> BARBERIS	99D	OMEG	450 $pp \rightarrow K^+K^-\pi^+\pi^-$
1710 ± 25		<sup>7</sup> FRENCH	99		300 $pp \rightarrow p_f(K^+K^-)p_S$
1707 ± 10		<sup>8</sup> AUGUSTIN	88	DM2	$J/\psi \rightarrow \gamma K^+K^-, K_S^0 K_S^0$
1698 ± 15		<sup>8</sup> AUGUSTIN	87	DM2	$J/\psi \rightarrow \gamma\pi^+\pi^-$
1720 ± 10 ± 10		<sup>9</sup> BALTRUSAIT.	.87	MRK3	$J/\psi \rightarrow \gamma K^+K^-$
1742 ± 15		<sup>8</sup> WILLIAMS	84	MP SF	200 $\pi^-N \rightarrow 2K_S^0 X$
1670 ± 50		BLOOM	83	CBAL	$J/\psi \rightarrow \gamma 2\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1744 ± 7 ± 5	381	<sup>10,11</sup> DOBBS	15		$J/\psi \rightarrow \gamma\pi^+\pi^-$
1705 ± 11 ± 5	237	<sup>10,11</sup> DOBBS	15		$\psi(2S) \rightarrow \gamma\pi^+\pi^-$
1706 ± 4 ± 5	1.0k	<sup>10,11</sup> DOBBS	15		$J/\psi \rightarrow \gamma K^+K^-$
1690 ± 8 ± 3	349	<sup>10,11</sup> DOBBS	15		$\psi(2S) \rightarrow \gamma K^+K^-$
1750 ± 13		AMSLER	06	CBAR	1.64 $\overline{p}p \rightarrow K^+K^-\pi^0$
1747 ± 5	80k	<sup>12,13</sup> UMAN	06	E835	5.2 $\overline{p}p \rightarrow \eta\eta\pi^0$
1776 ± 15		VLADIMIRSK..	06	SPEC	40 $\pi^-\rho \rightarrow K_S^0 K_S^0 n$
1790 <sup>+40</sup> <sub>−30</sub>		<sup>3</sup> ABLIKIM	05	BES2	$J/\psi \rightarrow \phi\pi^+\pi^-$
1670 ± 20		<sup>12</sup> BINON	05	GAMS	33 $\pi^-\rho \rightarrow \eta\eta n$
1726 ± 7	74	<sup>13</sup> CHEKANOV	04	ZEUS	$e\rho \rightarrow K_S^0 K_S^0 X$

See key on page 601

## Meson Particle Listings

 $f_0(1710)$ 

1732±15	14	ANISOVICH	03	RVUE	
1682±16		TIKHOMIROV	03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
1670±26	3.6k	4,15	NICHITIU	02	OBLX
1770±12	16,17	ANISOVICH	99B	SPEC	0.6-1.2 $p\bar{p} \rightarrow \eta\eta\pi^0$
1730±15	4	BARBERIS	99	OMEG	450 $pp \rightarrow p_S p_f K^+ K^-$
1750±20	4	BARBERIS	99B	OMEG	450 $pp \rightarrow p_S p_f \pi^+ \pi^-$
1750±30	18	ANISOVICH	98B	RVUE	Compilation
1720±39		BAI	98H	BES	$J/\psi \rightarrow \gamma\pi^0\pi^0$
1775±1.5	57	19	BARKOV	98	$\pi^- p \rightarrow K_S^0 K_S^0 n$
1690±11		20	ABREU	96C	DLPH $Z^0 \rightarrow K^+ K^- + X$
1696±5	5	9	BAI	96C	BES $J/\psi \rightarrow \gamma K^+ K^-$
1781±8	8	4	BAI	96C	BES $J/\psi \rightarrow \gamma K^+ K^-$
1768±14		BALOSHIN	95	SPEC	40 $\pi^- C \rightarrow K_S^0 K_S^0 X$
1750±15	21	BUGG	95	MRK3	$J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
1620±16	9	BUGG	95	MRK3	$J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
1748±10	8	ARMSTRONG	93C	E760	$\bar{p}p \rightarrow \pi^0\eta\eta \rightarrow 6\gamma$
~1750		BREAKSTONE	93	SFM	$pp \rightarrow pp\pi^+\pi^-\pi^+\pi^-$
1744±15	22	ALDE	92D	GAM2	38 $\pi^- p \rightarrow \eta\eta n$
1713±10	23	ARMSTRONG	89D	OMEG	300 $pp \rightarrow ppK^+K^-$
1706±10	23	ARMSTRONG	89D	OMEG	300 $pp \rightarrow ppK_S^0 K_S^0$
1700±15	9	BOLONKIN	88	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
1720±60	4	BOLONKIN	88	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
1638±10	24	FALVARD	88	DM2	$J/\psi \rightarrow \phi K^+ K^-, K_S^0 K_S^0$
1690±4	25	FALVARD	88	DM2	$J/\psi \rightarrow \phi K^+ K^-, K_S^0 K_S^0$
1755±8	26	ALDE	86C	GAM2	38 $\pi^- p \rightarrow n2\eta$
1730±2	27	LONGACRE	86	RVUE	22 $\pi^- p \rightarrow n2K_S^0$
1650±50	28,29	BURKE	82	MRK2	$J/\psi \rightarrow \gamma2\rho$
1640±50		EDWARDS	82D	CBAL	$J/\psi \rightarrow \gamma2\eta$
1730±10 ±20	30	ETKIN	82C	MPS	23 $\pi^- p \rightarrow n2K_S^0$

1 From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.

2 In the SU(3) based model with a specific interference pattern of the  $f_2(1270)$ ,  $a_2^0(1320)$ ,

and  $f_2'(1525)$  mesons incoherently added to the  $f_0(1710)$  and non-resonant background. This state may be different from  $f_0(1710)$ , see CLOSE 05.

3  $J^P = 0^+$ .

4 T-matrix pole.

5 Supersedes BARBERIS 99 and BARBERIS 99B.

6  $J^P = 0^+$ , supersedes by ARMSTRONG 89D.

7  $N_B J^{PC}$  determination.

8  $J^P = 2^+$ .

9 Using CLEO-c data but not authored by the CLEO Collaboration.

10 From a fit to a Breit-Wigner line shape with fixed  $\Gamma = 135$  MeV.

11 Breit-Wigner mass.

12 Systematic errors not estimated.

13 K-matrix pole, assuming  $J^P = 0^+$ , from combined analysis of  $\pi^- p \rightarrow \pi^0\pi^0 n$ ,  $\pi^- p \rightarrow K\bar{K}n$ ,  $\pi^+\pi^- \rightarrow \pi^+\pi^-$ ,  $\bar{p}p \rightarrow \pi^0\pi^0\pi^0$ ,  $\pi^0\eta\eta$ ,  $\pi^0\pi^0\eta$ ,  $\pi^+\pi^-\pi^0$ ,  $K^+K^-\pi^0$ ,  $K_S^0 K_S^0\pi^0$ ,  $K^+K_S^0\pi^-$  at rest,  $\bar{p}n \rightarrow \pi^-\pi^-\pi^+$ ,  $K_S^0 K^-\pi^0$ ,  $K_S^0 K_S^0\pi^-$  at rest.

14 Decaying to  $f_0(1370)\pi\pi$ .

15  $J^P = 0^+$ .

16 Not seen by AMSLER 02.

17 T-matrix pole, assuming  $J^P = 0^+$ .

18 No  $J^{PC}$  determination.

19 No  $J^{PC}$  determination, width not determined.

20 From a fit to the  $0^+$  partial wave.

21 ALDE 92D combines all the GAMS-2000 data.

22  $J^P = 2^+$ , superseded by FRENCH 99.

23 From an analysis ignoring interference with  $f_2'(1525)$ .

24 From an analysis including interference with  $f_2'(1525)$ .

25 Superseded by ALDE 92D.

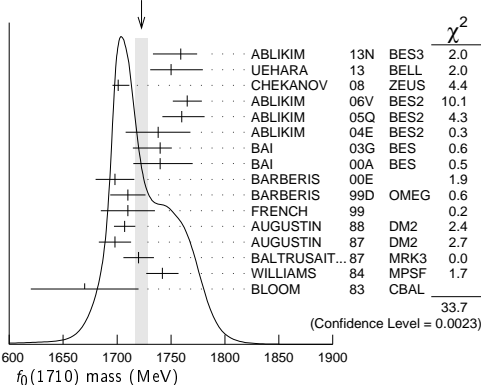
26 Uses MRK3 data. From a partial-wave analysis of data using a K-matrix formalism with 5 poles, but assuming spin 2. Fit with constrained inelasticity.

27  $J^P = 2^+$  preferred.

28 From fit neglecting nearby  $f_2'(1525)$ . Replaced by BLOOM 83.

29 Superseded by LONGACRE 86.

WEIGHTED AVERAGE  
1723±6-5 (Error scaled by 1.6)

 $f_0(1710)$  WIDTH

VALUE (MeV)		EVTS	DOCUMENT ID	TECN	COMMENT
<b>139 ± 8</b>	<b>OUR AVERAGE</b>		Error includes scale factor of 1.1.		
172 ± 10	+32 -16	5.5k	1	ABLIKIM 13N	BES3 $e^+e^- \rightarrow J/\psi \rightarrow \gamma\eta\eta$
139 ± 11	+96 -50			UEHARA 13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$
100 ± 24	+7 -22	4k	2	CHEKANOV 08	ZEUS $ep \rightarrow K_S^0 K_S^0 X$
145 ± 8	±69			ABLIKIM 06V	BES2 $e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^+\pi^-$
125 ± 25	+10 -15		3	ABLIKIM 05Q	BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^- K^+ K^-$
125 ± 20				ABLIKIM 04E	BES2 $J/\psi \rightarrow \omega K^+ K^-$
166 ± 5	+15 -10		4	BAI 03G	BES $J/\psi \rightarrow \gamma K\bar{K}$
120 ± 50	-40		4	BAI 00A	BES $J/\psi \rightarrow \gamma(\pi^+\pi^-\pi^+\pi^-)$
120 ± 26			5	BARBERIS 00E	450 $pp \rightarrow p_f \eta\eta p_S$
126 ± 16	±18		6	BARBERIS 99D	OMEG 450 $pp \rightarrow K^+ K^-, \pi^+\pi^-$
105 ± 34			7	FRENCH 99	300 $pp \rightarrow p_f(K^+ K^-) p_S$
166.4 ± 33.2			8	AUGUSTIN 88	DM2 $J/\psi \rightarrow \gamma K^+ K^-, K_S^0 K_S^0$
136 ± 28			8	AUGUSTIN 87	DM2 $J/\psi \rightarrow \gamma\pi^+\pi^-$
130 ± 20			9	BALTRUSAITIS 87	MRK3 $J/\psi \rightarrow \gamma K^+ K^-$
57 ± 38			10	WILLIAMS 84	MPSF 200 $\pi^- N \rightarrow 2K_S^0 X$
160 ± 80				BLOOM 83	CBAL $J/\psi \rightarrow \gamma 2\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
148 ± 40	-30			AMSLER 06	CBAR 1.64 $\bar{p}p \rightarrow K^+ K^-\pi^0$
188 ± 13		80k	3,11	UMAN 06	E835 5.2 $\bar{p}p \rightarrow \eta\eta\pi^0$
250 ± 30				VLADIMIRSKY 06	SPEC 40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
270 ± 60	-30		12	ABLIKIM 05	BES2 $J/\psi \rightarrow \phi\pi^+\pi^-$
260 ± 50			3	BINON 05	GAMS 33 $\pi^- p \rightarrow \eta\eta n$
38 ± 20	-14	74	11	CHEKANOV 04	ZEUS $ep \rightarrow K_S^0 K_S^0 X$
144 ± 30			13,14	ANISOVICH 03	RVUE
320 ± 50	-20		14,15	ANISOVICH 03	RVUE
102 ± 26				TIKHOMIROV 03	SPEC 40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
267 ± 44		3651	4,16	NICHITIU 02	OBLX
120 ± 40			17,18	ANISOVICH 99B	SPEC 0.6-1.2 $p\bar{p} \rightarrow \eta\eta\pi^0$
100 ± 25			4	BARBERIS 99	OMEG 450 $pp \rightarrow p_S p_f K^+ K^-$
160 ± 30			4	BARBERIS 99B	OMEG 450 $pp \rightarrow p_S p_f \pi^+\pi^-$
250 ± 140			19	ANISOVICH 98B	RVUE Compilation
30 ± 7		57	20	BARKOV 98	$\pi^- p \rightarrow K_S^0 K_S^0 n$
103 ± 18	+30 -11		9	BAI 96C	BES $J/\psi \rightarrow \gamma K^+ K^-$
85 ± 24	+22 -19		4	BAI 96C	BES $J/\psi \rightarrow \gamma K^+ K^-$
56 ± 19				BALOSHIN 95	SPEC 40 $\pi^- C \rightarrow K_S^0 K_S^0 X$
160 ± 40			21	BUGG 95	MRK3 $J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
160 ± 60	-20		9	BUGG 95	MRK3 $J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$
264 ± 25			8	ARMSTRONG 93C	E760 $\bar{p}p \rightarrow \pi^0\eta\eta \rightarrow 6\gamma$
200 to 300				BREAKSTONE 93	SFM $pp \rightarrow pp\pi^+\pi^-\pi^+\pi^-$
< 80 90% CL			22	ALDE 92D	GAM2 38 $\pi^- p \rightarrow \eta\eta N^*$
181 ± 30			23	ARMSTRONG 89D	OMEG 300 $pp \rightarrow ppK^+ K^-$
104 ± 30			23	ARMSTRONG 89D	OMEG 300 $pp \rightarrow ppK_S^0 K_S^0$
30 ± 20			9	BOLONKIN 88	SPEC 40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
350 ± 150			4	BOLONKIN 88	SPEC 40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
148 ± 17			24	FALVARD 88	DM2 $J/\psi \rightarrow \phi K^+ K^-, K_S^0 K_S^0$
184 ± 6			25	FALVARD 88	DM2 $J/\psi \rightarrow \phi K^+ K^-, K_S^0 K_S^0$
122 ± 74	-15		26	LONGACRE 86	RVUE 22 $\pi^- p \rightarrow n2K_S^0$
200 ± 100				BURKE 82	MRK2 $J/\psi \rightarrow \gamma 2\rho$
220 ± 100	-70		27,28	EDWARDS 82D	CBAL $J/\psi \rightarrow \gamma 2\eta$
200 ± 156	-9		29	ETKIN 82B	MPS 23 $\pi^- p \rightarrow n2K_S^0$

1 From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.

2 In the SU(3) based model with a specific interference pattern of the  $f_2(1270)$ ,  $a_2^0(1320)$ , and  $f_2'(1525)$  mesons incoherently added to the  $f_0(1710)$  and non-resonant background.

3 Breit-Wigner width.

4  $J^P = 0^+$ .

5 T-matrix pole.

6 Supersedes BARBERIS 99 and BARBERIS 99B.

7  $J^P = 0^+$ , supersedes by ARMSTRONG 89D.

8  $N_B J^{PC}$  determination.

9  $J^P = 2^+$ .

10 No  $J^{PC}$  determination.

11 Systematic errors not estimated.

12 This state may be different from  $f_0(1710)$ , see CLOSE 05.

13 (Solution I)

14 K-matrix pole, assuming  $J^P = 0^+$ , from combined analysis of  $\pi^- p \rightarrow \pi^0\pi^0 n$ ,  $\pi^- p \rightarrow K\bar{K}n$ ,  $\pi^+\pi^- \rightarrow \pi^+\pi^-$ ,  $\bar{p}p \rightarrow \pi^0\pi^0\pi^0$ ,  $\pi^0\eta\eta$ ,  $\pi^0\pi^0\eta$ ,  $\pi^+\pi^-\pi^0$ ,  $K^+K^-\pi^0$ ,  $K_S^0 K_S^0\pi^0$ ,  $K^+K_S^0\pi^-$  at rest,  $\bar{p}n \rightarrow \pi^-\pi^-\pi^+$ ,  $K_S^0 K^-\pi^0$ ,  $K_S^0 K_S^0\pi^-$  at rest.

# Meson Particle Listings

## $f_0(1710)$

- 15 (Solution I)
- 16 Decaying to  $f_0(1370)\pi\pi$ .
- 17  $J^P = 0^+$ .
- 18 Not seen by AMSLER 02.
- 19 T-matrix pole, assuming  $J^P = 0^+$
- 20 No  $J^{PC}$  determination.
- 21 From a fit to the  $0^+$  partial wave.
- 22 ALDE 92d combines all the GAMS-2000 data.
- 23  $J^P = 2^+$ , ( $0^+$  excluded).
- 24 From an analysis ignoring interference with  $f_2'(1525)$ .
- 25 From an analysis including interference with  $f_2'(1525)$ .
- 26 Uses MRK3 data. From a partial-wave analysis of data using a K-matrix formalism with 5 poles, but assuming spin 2. Fit with constrained inelasticity.
- 27  $J^P = 2^+$  preferred.
- 28 From fit neglecting nearby  $f_2'(1525)$ . Replaced by BLOOM 83.
- 29 From an amplitude analysis of the  $K_S^0 K_S^0$  system, superseded by LONGACRE 86.

### $f_0(1710)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\bar{K}$	seen
$\Gamma_2$ $\eta\eta$	seen
$\Gamma_3$ $\pi\pi$	seen
$\Gamma_4$ $\gamma\gamma$	
$\Gamma_5$ $\omega\omega$	seen

### $f_0(1710)$ $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$		$\Gamma_1\Gamma_4/\Gamma$			
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
$12^{+3+27}_{-2-8}$		UEHARA	13	BELL	$\gamma\gamma \rightarrow K_S^0 K_S^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<480	95	ALBRECHT	90G	ARG	$\gamma\gamma \rightarrow K^+ K^-$
<110	95	BEHREND	89C	CELL	$\gamma\gamma \rightarrow K_S^0 K_S^0$
<280	95	ALTHOFF	85B	TASS	$\gamma\gamma \rightarrow K\bar{K}\pi$
1 Assuming helicity 2.					
$\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$		$\Gamma_3\Gamma_4/\Gamma$			
VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT	
<0.82	95	1 BARATE	00E	ALEP	$\gamma\gamma \rightarrow \pi^+ \pi^-$
1 Assuming spin 0.					

### $f_0(1710)$ BRANCHING RATIOS

$\Gamma(K\bar{K})/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen	1004	<sup>1</sup> DOBBS	15	$J/\psi \rightarrow \gamma K^+ K^-$	
seen	349	<sup>1</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K^+ K^-$	
$0.36 \pm 0.12$		ALBALADEJO	08	RVUE	
$0.38^{+0.09}_{-0.19}$		<sup>2</sup> LONGACRE	86	MPS	$22 \pi^- p \rightarrow n 2 K_S^0$
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					
<sup>2</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles, but assuming spin 2. Fit with constrained inelasticity.					

$\Gamma(\eta\eta)/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$	
VALUE		DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.22 \pm 0.12$		ALBALADEJO 08	RVUE
$0.18^{+0.03}_{-0.13}$		<sup>1</sup> LONGACRE 86	RVUE
<sup>1</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles, but assuming spin 2. Fit with constrained inelasticity.			

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$		$\Gamma_3/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	381	<sup>1</sup> DOBBS 15		$J/\psi \rightarrow \gamma \pi^+ \pi^-$
seen	237	<sup>1</sup> DOBBS 15		$\psi(2S) \rightarrow \gamma \pi^+ \pi^-$
not seen		AMSLER 02	CBAR	$0.9 \bar{p}p \rightarrow \pi^0 \eta\eta, \pi^0 \pi^0 \pi^0$
$0.039^{+0.002}_{-0.024}$		<sup>2</sup> LONGACRE 86	RVUE	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.				
<sup>2</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles, but assuming spin 2. Fit with constrained inelasticity.				

$\Gamma(\pi\pi)/\Gamma(K\bar{K})$		$\Gamma_3/\Gamma_1$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$0.41^{+0.11}_{-0.17}$		ABLIKIM	06v	BES2	$e^+e^- \rightarrow J/\psi \rightarrow \gamma \pi^+ \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.32 \pm 0.14$		ALBALADEJO	08	RVUE	
< 0.11	95	1 ABLIKIM	04E	BES2	$J/\psi \rightarrow \omega K^+ K^-$
$5.8^{+9.1}_{-5.5}$		2 ANISOVICH	02D	SPEC	Combined fit
$0.2 \pm 0.024 \pm 0.036$		BARBERIS	99D	OMEG	$450\bar{p}p \rightarrow K^+ K^-, \pi^+ \pi^-$
$0.39 \pm 0.14$		ARMSTRONG	91	OMEG	$300\bar{p}p \rightarrow \bar{p}p\pi\pi, \bar{p}pK\bar{K}$
1 Using data from ABLIKIM 04A.					
2 From a combined K-matrix analysis of Crystal Barrel ( $0. \bar{p}\bar{p} \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta\eta, \pi^0 \pi^0 \eta$ ), GAMS ( $\pi p \rightarrow \pi^0 \pi^0 n, \eta\eta n, \eta\eta' n$ ), and BNL ( $\pi p \rightarrow K\bar{K} n$ ) data.					

$\Gamma(\eta\eta)/\Gamma(K\bar{K})$		$\Gamma_2/\Gamma_1$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$0.48 \pm 0.15$		BARBERIS	00E		$450\bar{p}p \rightarrow \bar{p}f\eta\eta\rho_S$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.46^{+0.70}_{-0.38}$		1 ANISOVICH	02D	SPEC	Combined fit
<0.02	90	2 PROKOSHKIN	91	GA24	$300\pi^- p \rightarrow \pi^- \rho\eta\eta$
1 From a combined K-matrix analysis of Crystal Barrel ( $0. \bar{p}\bar{p} \rightarrow \pi^0 \pi^0 \pi^0, \pi^0 \eta\eta, \pi^0 \pi^0 \eta$ ), GAMS ( $\pi p \rightarrow \pi^0 \pi^0 n, \eta\eta n, \eta\eta' n$ ), and BNL ( $\pi p \rightarrow K\bar{K} n$ ) data.					
2 Combining results of GAM4 with those of ARMSTRONG 89D.					

$\Gamma(\omega\omega)/\Gamma_{\text{total}}$		$\Gamma_5/\Gamma$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
seen	180	ABLIKIM	06H	BES	$J/\psi \rightarrow \gamma\omega\omega$

### $f_0(1710)$ REFERENCES

DOBBS 15 PR D91 052006 S. Dobbs *et al.* (NWES)

ABLIKIM 13N PR D87 092009 Ablikim M. *et al.* (BES III Collab.)

UEHARA 13 PTEP 2013 123C01 S. Uehara *et al.* (BELLE Collab.)

ALBALADEJO 08 PRL 101 252002 M. Albaladejo, J.A. Oller

CHEKANOV 08 PRL 101 112003 S. Chekanov *et al.* (ZEUS Collab.)

ABLIKIM 06H PR D73 112007 M. Ablikim *et al.* (BES Collab.)

ABLIKIM 06V PL B642 441 M. Ablikim *et al.* (BES Collab.)

AMSLER 06 PL B639 165 C. Amsler *et al.* (CBAR Collab.)

UMAN 06 PR D73 052009 I. Uman *et al.* (FNAL E835)

VLADIMIRSK... 06 PAN 69 493 V.V. Vladimirov *et al.* (ITEP, Moscow)

Translated from YAF 69 515.

ABLIKIM 05 PL B607 243 M. Ablikim *et al.* (BES Collab.)

ABLIKIM 05Q PR D72 092002 M. Ablikim *et al.* (BES Collab.)

BINON 05 PAN 68 960 F. Binon *et al.*

Translated from YAF 68 998.

CLOSE 05 PR D71 094022 F.E. Close, Q. Zhao

ABLIKIM 04A PL B598 149 M. Ablikim *et al.* (BES Collab.)

ABLIKIM 04E PL B603 138 M. Ablikim *et al.* (BES Collab.)

CHEKANOV 04 PL B578 33 S. Chekanov *et al.* (ZEUS Collab.)

PDG 04 PL B592 1 S. Eidelman *et al.* (PDG Collab.)

ANISOVICH 03 EPJ A16 229 V.V. Anisovich *et al.*

BAI 03G PR D68 052003 J.Z. Bai *et al.* (BES Collab.)

TIKHOMIROV 03 PAN 66 828 G.D. Tikhomirov *et al.*

Translated from YAF 66 860.

AMSLER 02 EPJ C23 29 C. Amsler *et al.*

ANISOVICH 02D PAN 65 1545 V.V. Anisovich *et al.*

Translated from YAF 65 1583.

NICHTIU 02 PL B545 261 F. Nichtiu *et al.* (OBELIX Collab.)

BAI 00A PL B472 207 J.Z. Bai *et al.* (BES Collab.)

BARATE 00E PL B472 189 R. Barate *et al.* (ALEPH Collab.)

BARBERIS 00E PL B479 59 D. Barberis *et al.* (WA 102 Collab.)

ANISOVICH 99B PL B449 154 A.V. Anisovich *et al.*

BARBERIS 99 PL B453 305 D. Barberis *et al.* (Omega Expt.)

BARBERIS 99B PL B453 316 D. Barberis *et al.* (Omega Expt.)

BARBERIS 99D PL B462 462 D. Barberis *et al.* (Omega Expt.)

FRENCH 99 PL B460 213 B. French *et al.* (WA76 Collab.)

ANISOVICH 98B SPU 41 419 V.V. Anisovich *et al.*

Translated from UFN 168 481.

BAI 98H PRL 81 1179 J.Z. Bai *et al.* (BES Collab.)

BARKOV 98 JET PL 68 764 B.P. Barkov *et al.*

ABREU 96C PL B379 309 P. Abreu *et al.* (DELPHI Collab.)

BAI 96C PRL 77 3959 J.Z. Bai *et al.* (BES Collab.)

BALOSHIN 95 PAN 58 46 O.N. Baloshin *et al.* (ITEP)

Translated from YAF 58 50.

BUGG 95 PL B353 378 D.V. Bugg *et al.* (LOQM, PNPI, WASH)

ARMSTRONG 93C PL B307 394 T.A. Armstrong *et al.* (FNAL, FERR, GENO+)

BREAKSTONE 93 ZPHY C58 251 A.M. Breakstone *et al.* (IOWA, CERN, DORT+)

ALDE 92D PL B284 457 D.M. Alde *et al.* (GAM2 Collab.)

SJNP 54 451 D.M. Alde *et al.* (GAM2 Collab.)

Translated from YAF 54 745.

ARMSTRONG 91 ZPHY C51 351 T.A. Armstrong *et al.* (ATHU, BARI, BIRM+)

PROKOSHKIN 91 SPD 36 155 Y.D. Prokoshkin (GAM2, GAM4 Collab.)

Translated from DANS 316 900.

ALBRECHT 90G ZPHY C48 183 H. Albrecht *et al.* (ARGUS Collab.)

ARMSTRONG 89D PL B227 186 T.A. Armstrong, M. Benayoun (ATHU, BARI, BIRM+)

BEHREND 89C ZPHY C43 91 H.J. Behrend *et al.* (CELLO Collab.)

AUGUSTIN 88 PRL 60 2238 J.E. Augustin *et al.* (DM2 Collab.)

BOLONKIN 88 NP B309 426 B.V. Bolonkin *et al.* (ITEP, SERP)

FALVARD 88 PR D38 2706 A. Falvard *et al.* (CLER, FRAS, LALO+)

AUGUSTIN 87 ZPHY C36 369 J.E. Augustin *et al.* (LALO, CLER, FRAS+)

BALTUSAITIS... 87 PR D35 2077 R.M. Baltusaitis *et al.* (Mark III Collab.)

ALDE 86C PL B182 105 D.M. Alde *et al.* (SERP, BELG, LANL, LAPP)

LONGACRE 86 PL B177 223 R.S. Longacre *et al.* (BNL, BRAN, CUNY+)

ALTHOFF 85B ZPHY C29 189 M. Althoff *et al.* (TASSO Collab.)

WILLIAMS 84 PR D30 877 E.G.H. Williams *et al.* (VAND, NDAM, TUFTS+)

BLOOM 83 ARNS 33 143 E.D. Bloom, C. Peck (SLAC, CIT)

BURKE 82 PRL 49 132 D.L. Burke *et al.* (LBL, SLAC)

EDWARDS 82D PRL 48 458 C. Edwards *et al.* (CIT, HARV, PRIN+)

ETKIN 82B PR D25 1786 A. Etkin *et al.* (BNL, CUNY, TUFTS, VAND)

ETKIN 82C PR D25 2446 A. Etkin *et al.* (BNL, CUNY, TUFTS, VAND)



See key on page 601

## Meson Particle Listings

 $\eta(1760), \pi(1800)$  $\eta(1760)$ 

$$J^G(J^{PC}) = 0^+(0^-+)$$

OMITTED FROM SUMMARY TABLE

Seen by DM2 in the  $\rho\rho$  system (BISELLO 89B). Structure in this region has been reported before in the same system (BALTRUSAITIS 86B) and in the  $\omega\omega$  system (BALTRUSAITIS 85C, BISELLO 87).

 $\eta(1760)$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1751±15 OUR AVERAGE</b>				
1768 <sup>+24</sup> <sub>-25</sub> ±10	465	<sup>1</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$
1744±10±15	1045	<sup>2</sup> ABLIKIM	06H BES	$J/\psi \rightarrow \gamma\omega\omega$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1703 <sup>+12</sup> <sub>-11</sub> ±2		<sup>3</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$
1760±11	320	<sup>4</sup> BISELLO	89B DM2	$J/\psi \rightarrow 4\pi\gamma$

<sup>1</sup> From a single-resonance fit.<sup>2</sup> From a partial wave analysis including  $\eta(1760)$ ,  $f_0(1710)$ ,  $f_2(1640)$ , and  $f_2(1910)$ .<sup>3</sup> From a two-resonance fit.<sup>4</sup> Estimated by us from various fits. Systematic uncertainties not estimated. $\eta(1760)$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>240±30 OUR AVERAGE</b>				
224 <sup>+62</sup> <sub>-56</sub> ±25	465	<sup>5</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$
244 <sup>+24</sup> <sub>-21</sub> ±25	1045	<sup>6</sup> ABLIKIM	06H BES	$J/\psi \rightarrow \gamma\omega\omega$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
42 <sup>+36</sup> <sub>-22</sub> ±15		<sup>7</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$
60±16	320	<sup>8</sup> BISELLO	89B DM2	$J/\psi \rightarrow 4\pi\gamma$

<sup>5</sup> From a single-resonance fit.<sup>6</sup> From a partial wave analysis including  $\eta(1760)$ ,  $f_0(1710)$ ,  $f_2(1640)$ , and  $f_2(1910)$ .<sup>7</sup> From a two-resonance fit.<sup>8</sup> Estimated by us from various fits. Systematic uncertainties not estimated. $\eta(1760)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $4\pi$	
$\Gamma_2$ $2\pi^+2\pi^-$	seen
$\Gamma_3$ $\pi^+\pi^-2\pi^0$	seen
$\Gamma_4$ $\rho^0\rho^0$	seen
$\Gamma_5$ $\rho^+\rho^-$	seen
$\Gamma_6$ $2(\pi^+\pi^-\pi^0)$	
$\Gamma_7$ $\omega\omega$	seen
$\Gamma_8$ $\eta'\pi^+\pi^-$	seen
$\Gamma_9$ $\gamma\gamma$	seen

 $\eta(1760)$   $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

$\Gamma(\eta'\pi^+\pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	VALUE (eV)	EVTs	DOCUMENT ID	TECN	COMMENT	$\Gamma_8\Gamma_9/\Gamma$
<b>28.2<sup>+7.9</sup><sub>-7.5</sub>±3.7</b>		465	<sup>9</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
3.0 <sup>+2.0</sup> <sub>-1.2</sub> ±0.8		52	<sup>10</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$	
18 <sup>+13</sup> <sub>-10</sub> ±5		315	<sup>11</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^- \eta' \pi^+ \pi^-$	

<sup>9</sup> From a single-resonance fit.<sup>10</sup> From a two-resonance fit. For constructive interference with the X(1835).<sup>11</sup> From a two-resonance fit. For destructive interference with the X(1835). $\eta(1760)$  BRANCHING RATIOS

$\Gamma(2\pi^+2\pi^-)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
seen		BISELLO	89B DM2	$J/\psi \rightarrow \gamma 2\pi^+2\pi^-$	
$\Gamma(\pi^+\pi^-2\pi^0)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
seen		BISELLO	89B DM2	$J/\psi \rightarrow \gamma \pi^+\pi^-2\pi^0$	

 $\Gamma(\rho^0\rho^0)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
seen	BISELLO	89B DM2	$J/\psi \rightarrow \gamma \rho^0 \rho^0$	
seen	BALTRUSAITIS...86	MRK3	$J/\psi \rightarrow \gamma \rho^0 \rho^0$	

 $\Gamma(\rho^+\rho^-)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
seen	BISELLO	89B DM2	$J/\psi \rightarrow \gamma \rho^+ \rho^-$	
seen	BALTRUSAITIS...86	MRK3	$J/\psi \rightarrow \gamma \rho^+ \rho^-$	

 $\Gamma(\omega\omega)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
seen	BISELLO	87 DM2	$J/\psi \rightarrow \omega\omega$	
seen	BALTRUSAITIS...85C	MRK3	$J/\psi \rightarrow \gamma\omega\omega$	

 $\eta(1760)$  REFERENCES

ZHANG	12A	PR D86 052002	C.C. Zhang <i>et al.</i>	(BELLE Collab.)
ABLIKIM	06H	PR D73 112007	M. Ablikim <i>et al.</i>	(BES Collab.)
BISELLO	89B	PR D39 701	G. Busetto <i>et al.</i>	(DM2 Collab.)
BISELLO	87	PL B192 239	D. Bisello <i>et al.</i>	(PADO, CLER, FRAS+)
BALTRUSAITIS...	86	PR D33 629	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BALTRUSAITIS...	86B	PR D33 1222	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BALTRUSAITIS...	85C	PRL 55 1723	R.M. Baltrusaitis <i>et al.</i>	(CIT, UCS C+)

 $\pi(1800)$ 

$$J^G(J^{PC}) = 1^-(0^-+)$$

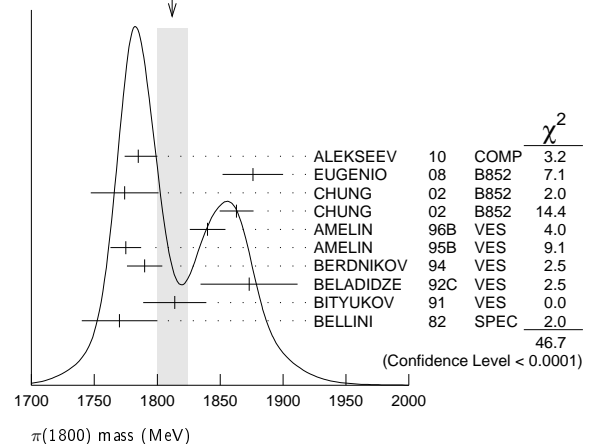
See also minireview under non- $q\bar{q}$  candidates in PDG 06, Journal of Physics **G33** 1 (2006).

 $\pi(1800)$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b>1812±12 OUR AVERAGE</b>					
Error includes scale factor of 2.3. See the ideogram below.					
1785±9 <sup>+12</sup> <sub>-6</sub>	420k	ALEKSEEV	10	COMP	190 $\pi^-\pi^-\pi^+\pi^+\pi^0$
1876±18±16	4k	<sup>1</sup> EUGENIO	08	B852	18 $\pi^-\pi^-\pi^+\pi^+\pi^0$
1774±18±20		<sup>2</sup> CHUNG	02	B852	18.3 $\pi^-\pi^-\pi^+\pi^+\pi^0$
					$\pi^+\pi^-\pi^-\pi^+\pi^0$
1863±9±10		<sup>3</sup> CHUNG	02	B852	18.3 $\pi^-\pi^-\pi^+\pi^+\pi^0$
					$\pi^+\pi^-\pi^-\pi^+\pi^0$
1840±10±10	1200	AMELIN	96B	VES	37 $\pi^-\pi^-\pi^+\pi^+\pi^0$
1775±7±10		<sup>4</sup> AMELIN	95B	VES	36 $\pi^-\pi^-\pi^+\pi^+\pi^0$
1790±14		<sup>5</sup> BERDNIKOV	94	VES	37 $\pi^-\pi^-\pi^+\pi^+\pi^0$
					$K^+K^-\pi^-\pi^+\pi^0$
1873±33±20		BELADIDZE	92C	VES	36 $\pi^-\pi^-\pi^+\pi^+\pi^0$
1814±10±23	426±57	BITYUKOV	91	VES	36 $\pi^-\pi^-\pi^+\pi^+\pi^0$
1770±30	1100	BELLINI	82	SPEC	40 $\pi^-\pi^-\pi^+\pi^+\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1737±5±15		AMELIN	99	VES	37 $\pi^-\pi^-\pi^+\pi^+\pi^0$

<sup>1</sup> From a single-pole fit.<sup>2</sup> In the  $f_0(980)\pi$  wave.<sup>3</sup> In the  $f_0(500)\pi$  wave.<sup>4</sup> From a fit to  $J^{PC} = 0^-+ f_0(980)\pi$ ,  $f_0(1370)\pi$  waves.<sup>5</sup> From a fit to  $J^{PC} = 0^-+ K_0^*(1430)K^-$  and  $f_0(980)\pi^-$  waves.

WEIGHTED AVERAGE  
1812±12 (Error scaled by 2.3)



# Meson Particle Listings

## $\pi(1800)$ , $f_2(1810)$

$\pi(1800)$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>208±12 OUR AVERAGE</b>					
208±22 <sup>+21</sup> <sub>-37</sub>	420k	ALEKSEEV	10	COMP	190 $\pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
221±26±38	4k	<sup>6</sup> EUGENIO	08	B852	— 18 $\pi^- p \rightarrow \eta \eta \pi^- p$
223±48±50		<sup>7</sup> CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
191±21±20		<sup>8</sup> CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
210±30±30	1200	AMELIN	96B	VES	— 37 $\pi^- A \rightarrow \eta \eta \pi^- A$
190±15±15		<sup>9</sup> AMELIN	95B	VES	— 36 $\pi^- A \rightarrow \pi^+ \pi^- \pi^- A$
210±70		<sup>10</sup> BERDNIKOV	94	VES	— 37 $\pi^- A \rightarrow K^+ K^- \pi^- A$
225±35±20		BELADIDZE	92c	VES	— 36 $\pi^- Be \rightarrow \pi^- \eta' \eta Be$
205±18±32	426±57	BITYUKOV	91	VES	— 36 $\pi^- C \rightarrow \pi^- \eta \eta C$
310±50	1100	BELLINI	82	SPEC	— 40 $\pi^- A \rightarrow 3\pi A$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
259±19±6		AMELIN	99	VES	37 $\pi^- A \rightarrow \omega \pi^- \pi^0 A^*$
<sup>6</sup> From a single-pole fit. <sup>7</sup> In the $f_0(980) \pi$ wave. <sup>8</sup> In the $f_0(500) \pi$ wave. <sup>9</sup> From a fit to $J^{PC} = 0^- + f_0(980) \pi$ , $f_0(1370) \pi$ waves. <sup>10</sup> From a fit to $J^{PC} = 0^- + K_0^0(1430) K^-$ and $f_0(980) \pi^-$ waves.					

$\pi(1800)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \pi^+ \pi^- \pi^-$	seen	
$\Gamma_2 f_0(500) \pi^-$	seen	
$\Gamma_3 f_0(980) \pi^-$	seen	
$\Gamma_4 f_0(1370) \pi^-$	seen	
$\Gamma_5 f_0(1500) \pi^-$	not seen	
$\Gamma_6 \rho \pi^-$	not seen	
$\Gamma_7 \eta \eta \pi^-$	seen	
$\Gamma_8 a_0(980) \eta$	seen	
$\Gamma_9 a_2(1320) \eta$	not seen	
$\Gamma_{10} f_2(1270) \pi$	not seen	
$\Gamma_{11} f_0(1370) \pi^-$	not seen	
$\Gamma_{12} f_0(1500) \pi^-$	seen	
$\Gamma_{13} \eta \eta' (958) \pi^-$	seen	
$\Gamma_{14} K_0^*(1430) K^-$	seen	
$\Gamma_{15} K^*(892) K^-$	not seen	

$\pi(1800)$ BRANCHING RATIOS				
$\Gamma(f_0(980) \pi^-)/\Gamma(f_0(500) \pi^-)$	$\Gamma_3/\Gamma_2$			
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.44±0.08±0.38</b>	<sup>11</sup> CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$

$\Gamma(f_0(980) \pi^-)/\Gamma(f_0(1370) \pi^-)$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.7±1.3	<sup>12</sup> AMELIN	95B	VES	— 36 $\pi^- A \rightarrow \pi^+ \pi^- \pi^- A$

$\Gamma(f_0(1370) \pi^-)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	BELLINI	82	SPEC	— 40 $\pi^- A \rightarrow 3\pi A$

$\Gamma(f_0(1500) \pi^-)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
not seen	CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$

$\Gamma(\rho \pi^-)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
not seen	BELLINI	82	SPEC	— 40 $\pi^- A \rightarrow 3\pi A$

$\Gamma(\rho\pi^-)/\Gamma(f_0(980)\pi^-)$					$\Gamma_6/\Gamma_3$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.25		CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
<0.14	90	AMELIN	95B	VES	— 36 $\pi^- A \rightarrow \pi^+ \pi^- \pi^- A$

$\Gamma(\eta\eta\pi^-)/\Gamma(\pi^+\pi^-\pi^-)$					$\Gamma_7/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.5 \pm 0.1$	1200	<sup>12</sup> AMELIN	96B VES	—	$37 \pi^- A \rightarrow \eta\eta\pi^- A$

$\Gamma(a_2(1320) \eta)/\Gamma_{\text{total}}$			
VALUE	DOCUMENT ID	TECN	COMMENT
not seen	EUGENIO	08	B852 18 $\pi^- p \rightarrow \eta \eta \pi^- p$

$\Gamma(f_2(1270) \pi)/\Gamma_{\text{total}}$			
VALUE	DOCUMENT ID	TECN	COMMENT
not seen	EUGENIO	08	B852 18 $\pi^- p \rightarrow \eta \eta \pi^- p$

$\Gamma(f_0(1370) \pi^-)/\Gamma_{\text{total}}$			
VALUE	DOCUMENT ID	TECN	COMMENT
not seen	EUGENIO	08	B852 18 $\pi^- p \rightarrow \eta \eta \pi^- p$

$\Gamma(f_0(1500) \pi^-)/\Gamma(a_0(980) \eta)$					
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.48 ±0.17	4k	<sup>12,13</sup> EUGENIO	08	B852	— 18 $\pi^- p \rightarrow \eta \eta \pi^- p$
0.030 <sup>+0.014</sup> <sub>-0.011</sub>		<sup>12</sup> ANISOVICH	01B	SPEC	0 0.6-1.94 $\rho \bar{p} \rightarrow \eta \eta \pi^0 \pi^0$
0.08 ±0.03	1200	<sup>12,14</sup> AMELIN	96B	VES	— 37 $\pi^- A \rightarrow \eta \eta \pi^- A$

$\Gamma(\eta \eta' (958) \pi^-)/\Gamma(\eta \eta \pi^-)$					
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.29±0.07		<sup>12</sup> BELADIDZE	92c	VES	— 36 $\pi^- Be \rightarrow \pi^- \eta' \eta Be$
0.3 ±0.1	426±57	<sup>12</sup> BITYUKOV	91	VES	— 36 $\pi^- C \rightarrow \pi^- \eta \eta C$

$\Gamma(K_0^*(1430) K^-)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	BERDNIKOV	94	VES	— 37 $\pi^- A \rightarrow K^+ K^- \pi^- A$

$\Gamma(K^*(892) K^-)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
not seen	BERDNIKOV	94	VES	— 37 $\pi^- A \rightarrow K^+ K^- \pi^- A$
<sup>11</sup> Assuming that $f_0(980)$ decays only to $\pi \pi$ . <sup>12</sup> Systematic errors not estimated. <sup>13</sup> From a single-pole fit. <sup>14</sup> Assuming that $f_0(1500)$ decays only to $\eta \eta$ and $a_0(980)$ decays only to $\eta \pi$ .				

$\pi(1800)$ REFERENCES					
ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i>	(COMPASS Collab.)	
EUGENIO	08	PL B660 466	P. Eugenio <i>et al.</i>	(BNL E852 Collab.)	
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)	
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)	
ANISOVICH	01B	PL B500 222	A.V. Anisovich <i>et al.</i>		
AMELIN	99	PAN 62 445	D.V. Amelin <i>et al.</i>	(VES Collab.)	
AMELIN	96B	Translated from YAF 62 487.	D.V. Amelin <i>et al.</i>	(SERP, TBIL)IGJPC	
AMELIN	95B	PAN 59 976	D.V. Amelin <i>et al.</i>	(SERP, TBIL)	
BERDNIKOV	94	PL B337 219	E.B. Berdnikov <i>et al.</i>	(SERP, TBIL)	
BELADIDZE	92C	SJNP 55 1535	G.M. Beladidze, S.I. Bityukov, G.V. Borisov	(SERP+)	
BITYUKOV	91	Translated from YAF 55 2748.	S.I. Bityukov <i>et al.</i>	(SERP, TBIL)	
BELLINI	82	PL B268 137	G. Bellini <i>et al.</i>	(MILA, BGNA, JINR)	

<div><math>f_2(1810)</math></div>	$I^G(J^{PC}) = 0^+(2^{++})$
OMITTED FROM SUMMARY TABLE	
Needs confirmation.	

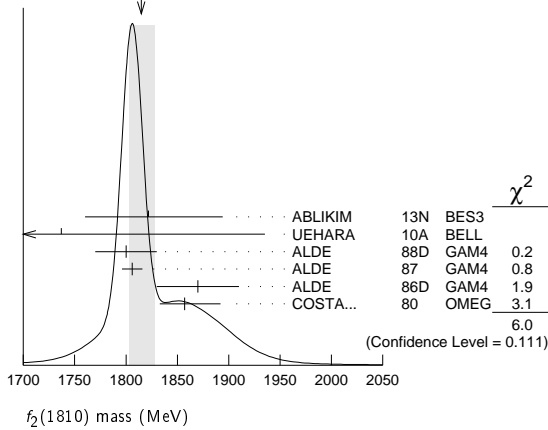
$f_2(1810)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1815±12 OUR AVERAGE</b>	Error includes scale factor of 1.4. See the ideogram below.			
1822 <sup>+29</sup> <sub>-24</sub> <sup>+66</sup> <sub>-57</sub>	5.5k	<sup>1</sup> ABLIKIM	13N	BES3 $e^+ e^- \rightarrow J/\psi \rightarrow \gamma \eta \eta$
1737±9 <sup>+198</sup> <sub>-65</sub>		<sup>2</sup> UEHARA	10A	BELL 10.6 $e^+ e^- \rightarrow e^+ e^- \eta \eta$
1800±30	40	ALDE	88D	GAM4 300 $\pi^- p \rightarrow \pi^- p 4\pi^0$
1806±10	1600	ALDE	87	GAM4 100 $\pi^- p \rightarrow 4\pi^0 n$
1870±40		<sup>3</sup> ALDE	86D	GAM4 100 $\pi^- p \rightarrow \eta \eta n$
1857 <sup>+35</sup> <sub>-24</sub>		<sup>4</sup> COSTA...	80	OMEG 10 $\pi^- p \rightarrow K^+ K^- n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1858 <sup>+18</sup> <sub>-71</sub>		<sup>5</sup> LONGACRE	86	RVUE Compilation
1799±15		<sup>6</sup> CASON	82	STRC 8 $\pi^+ p \rightarrow \Delta^{++} \pi^0 \pi^0$

<sup>1</sup> From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.  
<sup>2</sup> Breit-Wigner mass.  
<sup>3</sup> Seen in only one solution.  
<sup>4</sup> Error increased by spread of two solutions. Included in LONGACRE 86 global analysis.  
<sup>5</sup> From a partial-wave analysis of data using a K-matrix formalism with 5 poles. Includes compilation of several other experiments.  
<sup>6</sup> From an amplitude analysis of the reaction  $\pi^+ \pi^- \rightarrow 2\pi^0$ . The resonance in the  $2\pi^0$  final state is not confirmed by PROKOSHKIN 97.

See key on page 601

# Meson Particle Listings

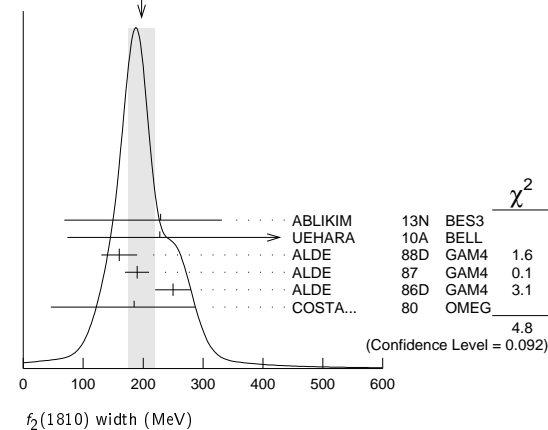
## $f_2(1810)$ , $X(1835)$

WEIGHTED AVERAGE  
1815±12 (Error scaled by 1.4)

### $f_2(1810)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>197±22 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.		
229 <sup>+52</sup> <sub>-42</sub>	88	7 ABLIKIM	13N BES3	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\eta\eta$
228 <sup>+21</sup> <sub>-20</sub>	234	8 UEHARA	10A BELL	$10.6 e^+e^- \rightarrow e^+e^- \eta\eta$
160±30	40	ALDE 88D	GAM4	$300 \pi^- p \rightarrow \pi^- p 4\pi^0$
190±20	1600	ALDE 87	GAM4	$100 \pi^- p \rightarrow 4\pi^0 n$
250±30		9 ALDE 86D	GAM4	$100 \pi^- p \rightarrow \eta\eta n$
185 <sup>+102</sup> <sub>-139</sub>		10 COSTA...	80 OMEG	$10 \pi^- p \rightarrow K^+ K^- n$
388 <sup>+15</sup> <sub>-21</sub>		11 LONGACRE 86	RVUE	Compilation
280 <sup>+42</sup> <sub>-35</sub>		12 CASON 82	STRC	$8 \pi^+ p \rightarrow \Delta^{++} \pi^0 \pi^0$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- 7 From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.
- 8 Breit-Wigner width.
- 9 Seen in only one solution.
- 10 Error increased by spread of two solutions. Included in LONGACRE 86 global analysis.
- 11 From a partial-wave analysis of data using a K-matrix formalism with 5 poles. Includes compilation of several other experiments.
- 12 From an amplitude analysis of the reaction  $\pi^+\pi^- \rightarrow 2\pi^0$ . The resonance in the  $2\pi^0$  final state is not confirmed by PROKOSHKIN 97.

WEIGHTED AVERAGE  
197±22 (Error scaled by 1.5)

### $f_2(1810)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \pi\pi$	
$\Gamma_2 \eta\eta$	seen
$\Gamma_3 4\pi^0$	seen
$\Gamma_4 K^+K^-$	
$\Gamma_5 \gamma\gamma$	seen

### $f_2(1810) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b>5.2±0.9+37.3-0.8-4.5</b>	13 UEHARA	10A BELL	$10.6 e^+e^- \rightarrow e^+e^- \eta\eta$

13 Including interference with the  $f_2'(1525)$  (parameters fixed to the values from the 2008 edition of this review, PDG 08) and  $f_2(1270)$ . May also be the  $f_0(1500)$ .

### $f_2(1810)$ BRANCHING RATIOS

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	AMSLER 02	CBAR	$0.9 \bar{p}p \rightarrow \pi^0 \eta\eta, \pi^0 \pi^0 \pi^0$
not seen	PROKOSHKIN 97	GAM2	$38 \pi^- p \rightarrow \pi^0 \pi^0 n$
0.21 <sup>+0.02</sup> <sub>-0.03</sub>	14 LONGACRE 86	RVUE	Compilation
0.44±0.03	15 CASON 82	STRC	$8 \pi^+ p \rightarrow \Delta^{++} \pi^0 \pi^0$

14 From a partial-wave analysis of data using a K-matrix formalism with 5 poles. Includes compilation of several other experiments.

15 Included in LONGACRE 86 global analysis.

$\Gamma(\eta\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
seen	ABLIKIM 13N	BES3	PWA of $J/\psi \rightarrow \gamma\eta\eta$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.008 <sup>+0.028</sup> <sub>-0.003</sub>	16 LONGACRE 86	RVUE	Compilation
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16 From a partial-wave analysis of data using a K-matrix formalism with 5 poles. Includes compilation of several other experiments.

$\Gamma(\pi\pi)/\Gamma(4\pi^0)$	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.75	ALDE 87	GAM4	$100 \pi^- p \rightarrow 4\pi^0 n$

$\Gamma(4\pi^0)/\Gamma(\eta\eta)$	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.8±0.3	ALDE 87	GAM4	$100 \pi^- p \rightarrow 4\pi^0 n$

$\Gamma(K^+K^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.003 <sup>+0.019</sup> <sub>-0.002</sub>	17 LONGACRE 86	RVUE	Compilation
seen	COSTA... 80	OMEG	$10 \pi^- p \rightarrow K^+ K^- n$

17 From a partial-wave analysis of data using a K-matrix formalism with 5 poles. Includes compilation of several other experiments.

### $f_2(1810)$ REFERENCES

ABLIKIM 13N	PR D87 092009	Ablikim M. et al.	(BES III Collab.)
UEHARA 10A	PR D82 114031	S. Uehara et al.	(BELLE Collab.)
PDG 08	PL B667 1	C. Amsler et al.	(PDG Collab.)
AMSLER 02	EPJ C23 29	C. Amsler et al.	
PROKOSHKIN 97	SPD 42 117	Y.D. Prokoshkin et al.	(SERP)
	Translated from DANS 353 323.		
ALDE 88D	SJNP 47 810	D.M. Alde et al.	(SERP, BELG, LANL, LAPP+)
	Translated from YAF 47 1273.		
ALDE 87	PL B198 286	D.M. Alde et al.	(LANL, BRUX, SERP, LAPP)
ALDE 86D	NP B269 485	D.M. Alde et al.	(BELG, LAPP, SERP, CERN+)
LONGACRE 86	PL B177 223	R.S. Longacre et al.	(BNL, BRAN, CUNY+)
CASON 82	PRL 48 1316	N.M. Cason et al.	(NDAM, ANL)
COSTA... 80	NP B175 402	G. Costa de Beauregard et al.	(BARI, BONN+)

### $X(1835)$

$$I^G(J^{PC}) = ?^?(0^-+)$$

### OMITTED FROM SUMMARY TABLE

Could be a superposition of two states, one with small width appearing as threshold enhancement in  $p\bar{p}$ , the other one with a larger width, decaying into  $\pi^+\pi^-\eta'$  and  $K_S^0 K_S^0 \eta$ . For the former ABLIKIM 12D determine  $J^{PC} = 0^-+$ .

### $X(1835)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1835.8±4.0 OUR AVERAGE</b>				
1844 ± 9 <sup>+16</sup> <sub>-25</sub>		ABLIKIM 15T	BES3	$J/\psi \rightarrow \gamma K_S^0 K_S^0 \eta$
1836.5±3.0 <sup>+5.6</sup> <sub>-2.1</sub>	4265	1 ABLIKIM 11C	BES3	$J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$
1833.7±6.1±2.7	264	ABLIKIM 05R	BES2	$J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$

Meson Particle Listings

X(1835), X(1840)

• • • We do not use the following data for averages, fits, limits, etc. • • •

1832	$+^{19}_{-5}$	$\pm 26$		<sup>2</sup> ABLIKIM	12D	BES3	$J/\psi \rightarrow \gamma p \overline{p}$
1877.3	$\pm 6.3$	$\pm^{3.4}_{-7.4}$		<sup>3</sup> ABLIKIM	11J	BES3	$J/\psi \rightarrow \omega(\eta \pi^+ \pi^-)$
1837	$+^{10}_{-12}$	$\pm^9_{-7}$	231	<sup>4,5</sup> ALEXANDER	10	CLEO	$J/\psi \rightarrow \gamma p \overline{p}$
1831	$\pm 7$			<sup>5,6</sup> ABLIKIM	05R	BES2	$J/\psi \rightarrow \gamma p \overline{p}$
1859	$+^3_{-10}$	$\pm^5_{-25}$		<sup>5</sup> BAI	03F	BES2	$J/\psi \rightarrow \gamma p \overline{p}$

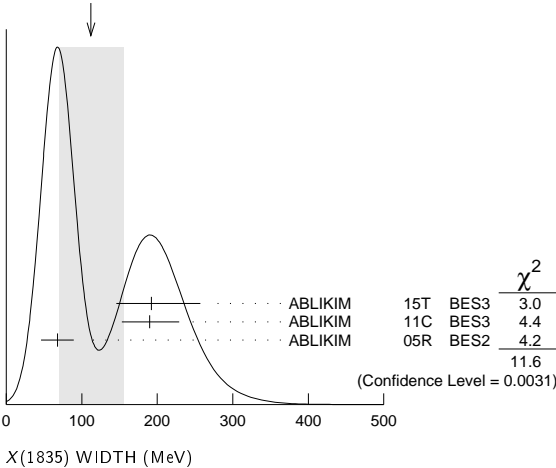
- <sup>1</sup> From a fit of the  $\pi^+ \pi^- \eta'$  mass distribution to a combination of  $\gamma f_1(1510)$ ,  $\gamma X(1835)$ , and two unconfirmed states  $\gamma X(2120)$ , and  $\gamma X(2370)$ , for  $M(p\overline{p}) < 2.8$  GeV, and accounting for backgrounds from non- $\eta'$  events and  $J/\psi \rightarrow \pi^0 \pi^+ \pi^- \eta'$ .
- <sup>2</sup> From the fit including final state interaction effects in isospin 0  $S$ -wave according to SIBIRTSEV 05A. Supersedes ABLIKIM 10G.
- <sup>3</sup> The selected process is  $J/\psi \rightarrow \omega \eta(980) \pi$ . This state may be due also to  $\eta_2(1870)$  or to a combination of  $X(1835)$  and  $\eta_2(1870)$ .
- <sup>4</sup> From a fit of the  $p\overline{p}$  mass distribution to a combination of  $\gamma X(1835)$ ,  $\gamma R$  with  $M(R) = 2100$  MeV and  $\Gamma(R) = 160$  MeV, and  $\gamma p\overline{p}$  phase space, for  $M(p\overline{p}) < 2.85$  GeV.
- <sup>5</sup> Evidence for a threshold enhancement in the  $p\overline{p}$  mass spectrum was also reported by ABE 02K, AUBERT,B 05L, and WANG 05A in  $B^+ \rightarrow p\overline{p} K^+$ , WANG 05A in  $B^0 \rightarrow p\overline{p} K_S^0$ , ABE 02W in  $\overline{B}^0 \rightarrow p\overline{p} D^0$ , DEL-AMO-SANCHEZ 12 in  $B \rightarrow D(D^*) p\overline{p}(\pi)$ , and WEI 08 in  $B^+ \rightarrow p\overline{p} \pi^+$  decays. Not seen by ATHAR 06 in  $\Upsilon(1S) \rightarrow p\overline{p} \gamma$ .
- <sup>6</sup> From the fit including final state interaction effects in isospin 0  $S$ -wave according to SIBIRTSEV 05A. Systematic errors not estimated.

X(1835) WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>112 ±40</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.			
192 +20 -17 +62 -43			ABLIKIM	15T BES3	$J/\psi \rightarrow \gamma K_S^0 K_S^0 \eta$
190 ± 9 +38 -36	4265	1	ABLIKIM	11C BES3	$J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$
67.7 ±20.3 ± 7.7	264		ABLIKIM	05R BES2	$J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 76	90	2	ABLIKIM	12D BES3	$J/\psi \rightarrow \gamma \rho \overline{\rho}$
57 ±12 +19 -4		3	ABLIKIM	11J BES3	$J/\psi \rightarrow \omega(\eta \pi^+ \pi^-)$
0 ±44 -0	231	4,5	ALEXANDER	10 CLEO	$J/\psi \rightarrow \gamma \rho \overline{\rho}$
< 153	90	5,6	ABLIKIM	05R BES2	$J/\psi \rightarrow \gamma \rho \overline{\rho}$
< 30		5	BAI	03F BES2	$J/\psi \rightarrow \gamma \rho \overline{\rho}$

- <sup>1</sup> From a fit of the  $\pi^+ \pi^- \eta'$  mass distribution to a combination of  $\gamma f_1(1510)$ ,  $\gamma X(1835)$ , and two unconfirmed states  $\gamma X(2120)$ , and  $\gamma X(2370)$ , for  $M(p\overline{p}) < 2.8$  GeV, and accounting for backgrounds from non- $\eta'$  events and  $J/\psi \rightarrow \pi^0 \pi^+ \pi^- \eta'$ .
- <sup>2</sup> From the fit including final state interaction effects in isospin 0  $S$ -wave according to SIBIRTSEV 05A. Supersedes ABLIKIM 10G.
- <sup>3</sup> The selected process is  $J/\psi \rightarrow \omega \eta(980) \pi$ . This state may be due also to  $\eta_2(1870)$  or to a combination of  $X(1835)$  and  $\eta_2(1870)$ .
- <sup>4</sup> From a fit of the  $p\overline{p}$  mass distribution to a combination of  $\gamma X(1835)$ ,  $\gamma R$  with  $M(R) = 2100$  MeV and  $\Gamma(R) = 160$  MeV, and  $\gamma p\overline{p}$  phase space, for  $M(p\overline{p}) < 2.85$  GeV.
- <sup>5</sup> Evidence for a threshold enhancement in the  $p\overline{p}$  mass spectrum was also reported by ABE 02K, AUBERT,B 05L, and WANG 05A in  $B^+ \rightarrow p\overline{p} K^+$ , WANG 05A in  $B^0 \rightarrow p\overline{p} K_S^0$ , ABE 02W in  $\overline{B}^0 \rightarrow p\overline{p} D^0$ , DEL-AMO-SANCHEZ 12 in  $B \rightarrow D(D^*) p\overline{p}(\pi)$ , and WEI 08 in  $B^+ \rightarrow p\overline{p} \pi^+$  decays. Not seen by ATHAR 06 in  $\Upsilon(1S) \rightarrow p\overline{p} \gamma$ .
- <sup>6</sup> From the fit including final state interaction effects in isospin 0  $S$ -wave according to SIBIRTSEV 05A. Systematic errors not estimated.

WEIGHTED AVERAGE  
112 $\pm$ 40 (Error scaled by 2.4)



X(1835) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $p \overline{p}$	seen
$\Gamma_2$ $\eta' \pi^+ \pi^-$	seen
$\Gamma_3$ $\gamma \gamma$	
$\Gamma_4$ $K_S^0 K_S^0 \eta$	seen

X(1835)  $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(\eta' \pi^+ \pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_2 \Gamma_3/\Gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
	<35.6	90	<sup>1</sup> ZHANG	12A	BELL $e^+ e^- \rightarrow e^+ e^- \eta' \pi^+ \pi^-$	
	<83	90	<sup>2</sup> ZHANG	12A	BELL $e^+ e^- \rightarrow e^+ e^- \eta' \pi^+ \pi^-$	
<sup>1</sup> From a two-resonance fit and constructive interference of the $\eta(1760)$ and $X(1835)$ , a significance of 2.8 $\sigma$ .						
<sup>2</sup> From a two-resonance fit and destructive interference of the $\eta(1760)$ and $X(1835)$ , a significance of 2.8 $\sigma$ .						

X(1835) BRANCHING RATIOS

$\Gamma(p\overline{p})/\Gamma(\eta' \pi^+ \pi^-)$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma_2$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
	0.333	ABLIKIM	05R	BES2 $J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$	

$\Gamma(\eta' \pi^+ \pi^-)/\Gamma(K_S^0 K_S^0 \eta)$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_4$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
	6.7 $\pm$ 1.8	<sup>1</sup> ABLIKIM	15T	BES3 $J/\psi \rightarrow \gamma K_S^0 K_S^0 \eta$	
<sup>1</sup> Using resutls from ABLIKIM 05R.					

X(1835) REFERENCES

ABLIKIM	15T	PRL 115 091803	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12D	PRL 108 112003	M. Ablikim <i>et al.</i>	(BES III Collab.) JPC
DEL-AMO-SA...	12	PR D85 092017	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
ZHANG	12A	PR D86 052002	C. C. Zhang <i>et al.</i>	(BELLE Collab.)
ABLIKIM	11C	PRL 106 072002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11J	PRL 107 182001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	10G	CPC 34 421	M. Ablikim <i>et al.</i>	(BES III Collab.)
ALEXANDER	10	PR D82 092002	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
WEI	08	PL B659 80	J.-T. Wei <i>et al.</i>	(BELLE Collab.)
ATHAR	06	PR D73 032001	S.B. Athar <i>et al.</i>	(CLEO Collab.)
ABLIKIM	05R	PRL 95 262001	M. Ablikim <i>et al.</i>	(BES Collab.)
AUBERT,B	05L	PR D72 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
SIBIRTSEV	05A	PR D71 054010	A. Sibirtsev, J. Haidenbauer	
WANG	05A	PL B617 141	M.-Z. Wang <i>et al.</i>	(BELLE Collab.)
BAI	03F	PRL 91 022001	J.Z. Bai <i>et al.</i>	(BES II Collab.)
ABE	02K	PRL 88 181803	K. Abe <i>et al.</i>	(BELLE Collab.)
ABE	02W	PRL 89 151802	K. Abe <i>et al.</i>	(BELLE Collab.)

X(1840)

$I^G(J^{PC}) = ?^?(?^{??})$

OMITTED FROM SUMMARY TABLE

X(1840) MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1842.2<math>\pm</math>4.2<math>+^{7.1}_{-2.6}</math></b>	0.6k	ABLIKIM	13U	BES3 $J/\psi \rightarrow \gamma 3(\pi^+ \pi^-)$

X(1840) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>83<math>\pm</math>14<math>\pm</math>11</b>	0.6k	ABLIKIM	13U	BES3 $J/\psi \rightarrow \gamma 3(\pi^+ \pi^-)$

X(1840) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $3(\pi^+ \pi^-)$	seen

X(1840) BRANCHING RATIOS

$\Gamma(3(\pi^+ \pi^-))/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
	seen	0.6k	ABLIKIM	13U	BES3 $J/\psi \rightarrow \gamma 3(\pi^+ \pi^-)$	

X(1840) REFERENCES

ABLIKIM	13U	PR D88 091502	M. Ablikim <i>et al.</i>	(BES III Collab.)
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See key on page 601

Meson Particle Listings

$a_1(1420)$ ,  $\phi_3(1850)$ ,  $\eta_2(1870)$

$a_1(1420)$

$I^G(J^{PC}) = 1^-(1^{++})$

OMITTED FROM SUMMARY TABLE

$a_1(1420)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
$1414^{+15}_{-13}$	<sup>1</sup> ADOLPH	15C	COMP	$190 \pi^- p \rightarrow \pi^- \pi^+ \pi^- p$
<sup>1</sup> Using the isobar model and partial-wave analysis with 88 waves.				

$a_1(1420)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
$153^{+8}_{-23}$	<sup>1</sup> ADOLPH	15C	COMP	$190 \pi^- p \rightarrow \pi^- \pi^+ \pi^- p$
<sup>1</sup> Using the isobar model and partial-wave analysis with 88 waves.				

$a_1(1420)$ DECAY MODES				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$f_0(980)\pi$	seen		

$a_1(1420)$ BRANCHING RATIOS				
$\Gamma(f_0(980)\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	<sup>1</sup> ADOLPH	15C	COMP	$190 \pi^- p \rightarrow \pi^- \pi^+ \pi^- p$
<sup>1</sup> Using the isobar model and partial-wave analysis with 88 waves.				

$a_1(1420)$ REFERENCES				
ADOLPH	15C	PRL 115 082001	C. Adolph <i>et al.</i>	(COMPASS Collab.)

$\phi_3(1850)$

$I^G(J^{PC}) = 0^-(3^{--})$

$\phi_3(1850)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$1854 \pm 7$ OUR AVERAGE				
$1855 \pm 10$		ASTON	88E	LASS $11 K^- p \rightarrow K^- K^+ \Lambda, K_S^0 K^\pm \pi^\mp \Lambda$
$1870^{+30}_{-20}$	430	ARMSTRONG	82	OMEG $18.5 K^- p \rightarrow K^- K^+ \Lambda$
$1850 \pm 10$	123	ALHARRAN	81B	HBC $8.25 K^- p \rightarrow K \bar{K} \Lambda$

$\phi_3(1850)$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$87^{+28}_{-23}$ OUR AVERAGE	Error includes scale factor of 1.2.			
$64 \pm 31$		ASTON	88E	LASS $11 K^- p \rightarrow K^- K^+ \Lambda, K_S^0 K^\pm \pi^\mp \Lambda$
$160^{+90}_{-50}$	430	ARMSTRONG	82	OMEG $18.5 K^- p \rightarrow K^- K^+ \Lambda$
$80^{+40}_{-30}$	123	ALHARRAN	81B	HBC $8.25 K^- p \rightarrow K \bar{K} \Lambda$

$\phi_3(1850)$ DECAY MODES				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$K \bar{K}$	seen		
$\Gamma_2$	$K \bar{K}^*(892) + \text{c.c.}$	seen		

$\phi_3(1850)$ BRANCHING RATIOS				
$\Gamma(K \bar{K}^*(892) + \text{c.c.})/\Gamma(K \bar{K})$				$\Gamma_2/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.55^{+0.85}_{-0.45}$	ASTON	88E	LASS	$11 K^- p \rightarrow K^- K^+ \Lambda, K_S^0 K^\pm \pi^\mp \Lambda$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.8 \pm 0.4$	ALHARRAN	81B	HBC	$8.25 K^- p \rightarrow K \bar{K} \pi \Lambda$

$\phi_3(1850)$ REFERENCES				
ASTON	88E	PL B208 324	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)IGJPC
ARMSTRONG	82	PL 110B 77	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)JP
ALHARRAN	81B	PL 101B 357	S. Al-Harran <i>et al.</i>	(BIRM, CERN, GLAS+)

$\eta_2(1870)$

$I^G(J^{PC}) = 0^+(2^{-+})$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

$\eta_2(1870)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$1842 \pm 8$ OUR AVERAGE				
$1835 \pm 12$		BARBERIS	00B	$450 p p \rightarrow p_f \eta \pi^+ \pi^- p_s$
$1844 \pm 13$		BARBERIS	00C	$450 p p \rightarrow p_f 4 \pi p_s$
$1840 \pm 25$		BARBERIS	97B	OMEG $450 p p \rightarrow p p 2(\pi^+ \pi^-)$
$1875 \pm 20 \pm 35$		ADOMEIT	96	CBAR $1.94 \bar{p} p \rightarrow \eta 3 \pi^0$
$1881 \pm 32 \pm 40$	26	KARCH	92	CBAL $e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1860 \pm 5 \pm 15$		ANISOVICH	00E	SPEC $0.9\text{--}1.94 \bar{p} p \rightarrow \eta 3 \pi^0$
$1840 \pm 15$		BAI	99	BES $J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$

$\eta_2(1870)$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$225 \pm 14$ OUR AVERAGE				
$235 \pm 22$		BARBERIS	00B	$450 p p \rightarrow p_f \eta \pi^+ \pi^- p_s$
$228 \pm 23$		BARBERIS	00C	$450 p p \rightarrow p_f 4 \pi p_s$
$200 \pm 40$		BARBERIS	97B	OMEG $450 p p \rightarrow p p 2(\pi^+ \pi^-)$
$200 \pm 25 \pm 45$		ADOMEIT	96	CBAR $1.94 \bar{p} p \rightarrow \eta 3 \pi^0$
$221 \pm 92 \pm 44$	26	KARCH	92	CBAL $e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$250 \pm 25^{+50}_{-35}$		ANISOVICH	00E	SPEC $0.9\text{--}1.94 \bar{p} p \rightarrow \eta 3 \pi^0$
$170 \pm 40$		BAI	99	BES $J/\psi \rightarrow \gamma \eta \pi^+ \pi^-$

$\eta_2(1870)$ DECAY MODES				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$\eta \pi \pi$			
$\Gamma_2$	$a_2(1320)\pi$			
$\Gamma_3$	$f_2(1270)\eta$			
$\Gamma_4$	$a_0(980)\pi$			
$\Gamma_5$	$\gamma \gamma$	seen		

$\eta_2(1870)$ BRANCHING RATIOS				
$\Gamma(a_2(1320)\pi)/\Gamma(f_2(1270)\eta)$				$\Gamma_2/\Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
$1.7 \pm 0.4$ OUR AVERAGE				
$1.60 \pm 0.40$	<sup>1</sup> ANISOVICH	11	SPEC	$0.9\text{--}1.94 p \bar{p}$
$20.4 \pm 6.6$	BARBERIS	00B		$450 p p \rightarrow p_f \eta \pi^+ \pi^- p_s$
$4.1 \pm 2.3$	ADOMEIT	96	CBAR	$1.94 \bar{p} p \rightarrow \eta 3 \pi^0$
<sup>1</sup> Reanalysis of ADOMEIT 96 and ANISOVICH 00E.				

$\Gamma(a_2(1320)\pi)/\Gamma(a_0(980)\pi)$				
VALUE	DOCUMENT ID	COMMENT		
$32.6 \pm 12.6$	BARBERIS	00B	$450 p p \rightarrow p_f \eta \pi^+ \pi^- p_s$	$\Gamma_2/\Gamma_4$
$\Gamma(a_0(980)\pi)/\Gamma(f_2(1270)\eta)$				$\Gamma_4/\Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.48 \pm 0.45$	<sup>2</sup> ANISOVICH	11	SPEC	$0.9\text{--}1.94 p \bar{p}$
<sup>2</sup> Reanalysis of ADOMEIT 96 and ANISOVICH 00E.				

$\Gamma(\gamma \gamma)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	KARCH	92	CBAL	$e^+ e^- \rightarrow e^+ e^- \eta \pi^0 \pi^0$

$\eta_2(1870)$ REFERENCES				
ANISOVICH	11	EPJ C71 1511	A.V. Anisovich <i>et al.</i>	(LOQM, RAL, PNPI)
ANISOVICH	00E	PL B477 19	A.V. Anisovich <i>et al.</i>	
BARBERIS	00B	PL B471 435	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00C	PL B471 440	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BAI	99	PL B446 356	J.Z. Bai <i>et al.</i>	(BES Collab.)
BARBERIS	97B	PL B413 217	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ADOMEIT	96	ZPHY C71 227	J. Adomeit <i>et al.</i>	(Crystal Barrel Collab.)
KARCH	92	ZPHY C54 33	K. Karch <i>et al.</i>	(Crystal Ball Collab.)

# Meson Particle Listings

$\pi_2(1880)$ ,  $\rho(1900)$ ,  $f_2(1910)$

$\pi_2(1880)$

$I^G(J^{PC}) = 1^-(2^--)$

$\pi(1880)$ MASS					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b>1895 ± 16 OUR AVERAGE</b>					
1929 ± 24 ± 18	4k	EUGENIO	08	B852	— 18 $\pi^- p \rightarrow \eta \eta \pi^- p$
1876 ± 11 ± 67	145k	LU	05	B852	— 18 $\pi^- p \rightarrow \omega \pi^- \pi^0 p$
2003 ± 88 ± 148	69k	KUHN	04	B852	— 18 $\pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$
1880 ± 20		ANISOVICH	01B	SPEC	0 0.6–1.94 $\bar{p} p \rightarrow \eta \eta \pi^0 \pi^0$

$\pi(1880)$ WIDTH					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b>235 ± 34 OUR AVERAGE</b>					
323 ± 87 ± 43	4k	EUGENIO	08	B852	— 18 $\pi^- p \rightarrow \eta \eta \pi^- p$
146 ± 17 ± 62	145k	LU	05	B852	— 18 $\pi^- p \rightarrow \omega \pi^- \pi^0 p$
306 ± 132 ± 121	69k	KUHN	04	B852	— 18 $\pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$
255 ± 45		ANISOVICH	01B	SPEC	0 0.6–1.94 $\bar{p} p \rightarrow \eta \eta \pi^0 \pi^0$

$\pi_2(1880)$ DECAY MODES	
Mode	
$\Gamma_1$	$\eta \eta \pi^-$
$\Gamma_2$	$a_0(980) \eta$
$\Gamma_3$	$a_2(1320) \eta$
$\Gamma_4$	$f_0(1500) \pi$
$\Gamma_5$	$f_1(1285) \pi$
$\Gamma_6$	$\omega \pi^- \pi^0$

$\Gamma(a_2(1320)\eta)/\Gamma(f_1(1285)\pi)$					$\Gamma_3/\Gamma_5$
VALUE	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
22.7 ± 7.3	69k	KUHN	04	B852	— 18 $\pi^- p \rightarrow \eta \pi^+ \pi^- \pi^- p$

$\Gamma(f_0(1500)\pi)/\Gamma(a_0(980)\eta)$					$\Gamma_4/\Gamma_2$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.26 <sup>+0.20</sup> <sub>−0.15</sub>		<sup>1</sup> ANISOVICH	01B	SPEC	0 0.6–1.94 $\bar{p} p \rightarrow \eta \eta \pi^0 \pi^0$
<sup>1</sup> Systematic errors not estimated.					

$\pi_2(1880)$ REFERENCES					
EUGENIO	08	PL B660 466	P. Eugenio <i>et al.</i>	(BNL E852 Collab.)	
LU	05	PRL 94 032002	M. Lu <i>et al.</i>	(BNL E852 Collab.)	
KUHN	04	PL B595 109	J. Kuhn <i>et al.</i>	(BNL E852 Collab.)	
ANISOVICH	01B	PL B500 222	A.V. Anisovich <i>et al.</i>		

$\rho(1900)$

$I^G(J^{PC}) = 1^+(1^{--})$

OMITTED FROM SUMMARY TABLE  
See our mini-review under the  $\rho(1700)$ .

$\rho(1900)$ MASS					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1909 ± 17 ± 25	54	<sup>1</sup> AUBERT	08s	BABR	10.6 $e^+ e^- \rightarrow \phi \pi^0 \gamma$
1880 ± 30		AUBERT	06D	BABR	10.6 $e^+ e^- \rightarrow 3\pi^+ 3\pi^- \gamma$
1860 ± 20		AUBERT	06D	BABR	10.6 $e^+ e^- \rightarrow 2(\pi^+ \pi^- \pi^0) \gamma$
1910 ± 10		<sup>2,3</sup> FRABETTI	04	E687	$\gamma p \rightarrow 3\pi^+ 3\pi^- p$
1870 ± 10		ANTONELLI	96	SPEC	$e^+ e^- \rightarrow$ hadrons
<sup>1</sup> From the fit with two resonances. <sup>2</sup> From a fit with two resonances with the JACOB 72 continuum. <sup>3</sup> Supersedes FRABETTI 01.					

$\rho(1900)$ WIDTH					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
48 ± 17 ± 2	54	<sup>4</sup> AUBERT	08s	BABR	10.6 $e^+ e^- \rightarrow \phi \pi^0 \gamma$
130 ± 30		AUBERT	06D	BABR	10.6 $e^+ e^- \rightarrow 3\pi^+ 3\pi^- \gamma$
160 ± 20		AUBERT	06D	BABR	10.6 $e^+ e^- \rightarrow 2(\pi^+ \pi^- \pi^0) \gamma$
37 ± 13		<sup>5,6</sup> FRABETTI	04	E687	$\gamma p \rightarrow 3\pi^+ 3\pi^- p$
10 ± 5		ANTONELLI	96	SPEC	$e^+ e^- \rightarrow$ hadrons
<sup>4</sup> From the fit with two resonances. <sup>5</sup> From a fit with two resonances with the JACOB 72 continuum. <sup>6</sup> Supersedes FRABETTI 01.					

$\rho(1900) \Gamma(i) \Gamma(e^+ e^-) / \Gamma^2(\text{total})$					$\Gamma_4/\Gamma \times \Gamma_6/\Gamma$
$\Gamma(\phi\pi)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	VALUE (units 10 <sup>−8</sup> )	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.2 ± 1.2 ± 0.8		54	<sup>7</sup> AUBERT	08s	BABR 10.6 $e^+ e^- \rightarrow \phi \pi^0 \gamma$
<sup>7</sup> From the fit with two resonances.					

$\rho(1900)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$6\pi$ seen
$\Gamma_2$	$3\pi^+ 3\pi^-$ seen
$\Gamma_3$	$2\pi^+ 2\pi^- 2\pi^0$
$\Gamma_4$	$\phi \pi$
$\Gamma_5$	hadrons seen
$\Gamma_6$	$e^+ e^-$ seen
$\Gamma_7$	$\bar{N} N$ not seen

$\rho(1900)$ BRANCHING RATIOS					$\Gamma_1/\Gamma$
$\Gamma(6\pi)/\Gamma_{\text{total}}$	VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
seen		8k	AKHMETSHIN 13	CMD3	$e^+ e^- \rightarrow 3\pi^+ 3\pi^-$
not seen			AGNELLO 02	OBLX	$\bar{\pi} p \rightarrow 3\pi^+ 2\pi^- \pi^0$
seen			FRABETTI 01	E687	$\gamma p \rightarrow 3\pi^+ 3\pi^- p$
seen			ANTONELLI 96	SPEC	$e^+ e^- \rightarrow$ hadrons

$\rho(1900)$ REFERENCES					
AKHMETSHIN 13	PL B723 82	R.R. Akhmetshin <i>et al.</i>	(CMD-3 Collab.)		
AUBERT 08s	PR D77 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)		
AUBERT 06D	PR D73 052003	B. Aubert <i>et al.</i>	(BABAR Collab.)		
FRABETTI 04	PL B578 290	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)		
AGNELLO 02	PL B527 39	M. Agnello <i>et al.</i>	(OBELIX Collab.)		
FRABETTI 01	PL B514 240	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)		
ANTONELLI 96	PL B365 427	A. Antonelli <i>et al.</i>	(FENICE Collab.)		
JACOB 72	PR D5 1847	M. Jacob, R. Slansky			

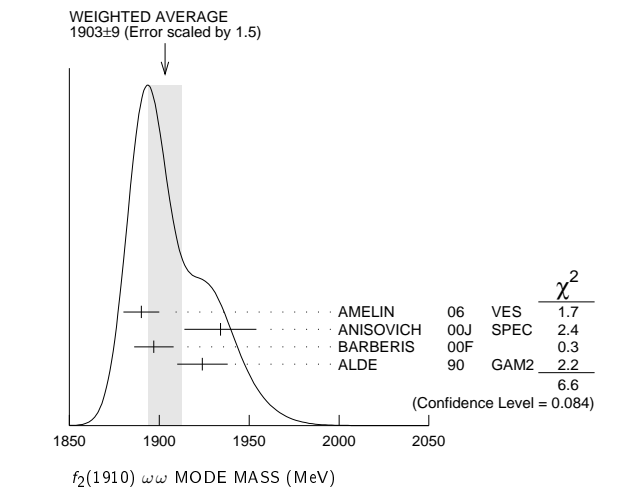
$f_2(1910)$

$I^G(J^{PC}) = 0^+(2^{++})$

OMITTED FROM SUMMARY TABLE  
We list here three different peaks with close masses and widths seen in the mass distributions of  $\omega\omega$ ,  $\eta\eta'$ , and  $K^+ K^-$  final states. ALDE 91B argues that they are of different nature.

$f_2(1910)$ MASS					

$f_2(1910) \omega\omega$ MODE					
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT		
<b>1903 ± 9 OUR AVERAGE</b> Error includes scale factor of 1.5. See the ideogram below.					
1890 ± 10	<sup>1</sup> AMELIN	06	VES	36 $\pi^- p \rightarrow \omega\omega \eta$	
1934 ± 20	ANISOVICH	00J	SPEC		
1897 ± 11	BARBERIS	00F		450 $p p \rightarrow p_f \omega\omega p_s$	
1924 ± 14	ALDE	90	GAM2	38 $\pi^- p \rightarrow \omega\omega \eta$	
<sup>1</sup> Supersedes BELADIDZE 92B.					



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# Meson Particle Listings

## $f_2(1910)$ , $a_0(1950)$

### $f_2(1910) \eta\eta'$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1934±16</b>	<sup>2</sup> BARBERIS	00A	450 $pp \rightarrow p_f \eta \eta' p_S$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1911±10	ALDE	91B	GAM2 38 $\pi^- p \rightarrow \eta \eta' n$
<sup>2</sup> Also compatible with $J^{PC}=1^-+$ .			

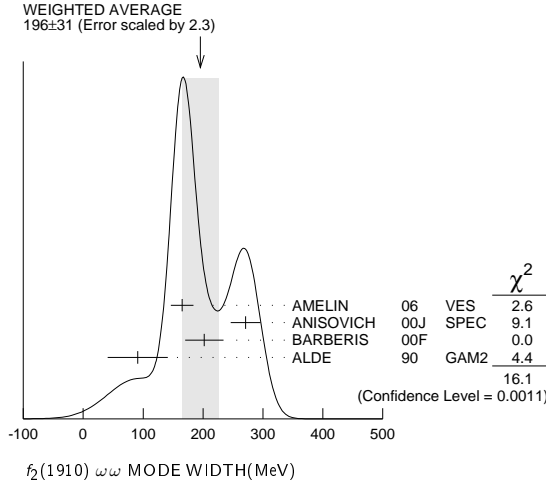
### $f_2(1910) K^+ K^-$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1941±18	AMSLER	06	CBAR 1.64 $\overline{p} p \rightarrow K^+ K^- \pi^0$

### $f_2(1910)$ WIDTH

### $f_2(1910) \omega\omega$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>196±31 OUR AVERAGE</b>	Error includes scale factor of 2.3. See the ideogram below.		
165±19	<sup>3</sup> AMELIN	06	VES 36 $\pi^- p \rightarrow \omega\omega n$
271±25	ANISOVICH	00J	SPEC
202±32	BARBERIS	00F	450 $pp \rightarrow p_f \omega\omega p_S$
91±5.0	ALDE	90	GAM2 38 $\pi^- p \rightarrow \omega\omega n$
<sup>3</sup> Supersedes BELADIDZE 92b.			



### $f_2(1910) \eta\eta'$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>141±41</b>	<sup>4</sup> BARBERIS	00A	450 $pp \rightarrow p_f \eta \eta' p_S$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
90±35	ALDE	91B	GAM2 38 $\pi^- p \rightarrow \eta \eta' n$
<sup>4</sup> Also compatible with $J^{PC}=1^-+$ .			

### $f_2(1910) K^+ K^-$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
120±40	AMSLER	06	CBAR 1.64 $\overline{p} p \rightarrow K^+ K^- \pi^0$

### $f_2(1910)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \pi^0 \pi^0$	
$\Gamma_2 K^+ K^-$	seen
$\Gamma_3 K_S^0 K_S^0$	
$\Gamma_4 \eta \eta$	seen
$\Gamma_5 \omega \omega$	seen
$\Gamma_6 \eta \eta'$	seen
$\Gamma_7 \eta' \eta'$	
$\Gamma_8 \rho \rho$	seen
$\Gamma_9 a_2(1320) \pi$	seen
$\Gamma_{10} f_2(1270) \eta$	seen

### $f_2(1910)$ BRANCHING RATIOS

$\Gamma(K^+ K^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
seen	AMSLER	06	CBAR 1.64 $\overline{p} p \rightarrow K^+ K^- \pi^0$

### $\Gamma(\pi^0 \pi^0)/\Gamma(\eta\eta')$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.1	ALDE	89	GAM2 38 $\pi^- p \rightarrow \eta \eta' n$

### $\Gamma(K_S^0 K_S^0)/\Gamma(\eta\eta')$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.066	90	BALOSHIN	86	SPEC 40 $\pi p \rightarrow K_S^0 K_S^0 n$

### $\Gamma(\eta\eta)/\Gamma(\eta\eta')$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.05	90	ALDE	91B	GAM2 38 $\pi^- p \rightarrow \eta \eta' n$

### $\Gamma(\omega\omega)/\Gamma(\eta\eta')$

VALUE	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
2.6±0.6	BARBERIS	00F 450 $pp \rightarrow p_f \omega\omega p_S$

### $\Gamma(\eta'\eta')/\Gamma_{\text{total}}$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
probably not seen	BARBERIS	00A	450 $pp \rightarrow p_f \eta' \eta' p_S$
possibly seen	BELADIDZE	92D	VES 37 $\pi^- p \rightarrow \eta' \eta' n$

### $\Gamma(\rho\rho)/\Gamma(\omega\omega)$

VALUE	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
2.6±0.4	BARBERIS	00F 450 $pp \rightarrow p_f \omega\omega p_S$

### $\Gamma(f_2(1270) \eta)/\Gamma(a_2(1320) \pi)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.09±0.05</b>	<sup>5</sup> ANISOVICH	11	SPEC 0.9–1.94 $p\overline{p}$
<sup>5</sup> Reanalysis of ADOMEIT 96 and ANISOVICH 00E.			

### $f_2(1910)$ REFERENCES

ANISOVICH	11	EPJ C71 1511	A.V. Anisovich et al.	(LOQM, RAL, PNPI)
AMELIN	06	PAN 69 690	D.V. Amelin et al.	(VES Collab.)
		Translated from YAF 69 715.		
AMSLER	06	PL B639 165	C. Amisler et al.	(CBAR Collab.)
ANISOVICH	00E	PL B477 19	A.V. Anisovich et al.	
ANISOVICH	00J	PL B491 47	A.V. Anisovich et al.	
BARBERIS	00A	PL B471 429	D. Barberis et al.	(WA 102 Collab.)
BARBERIS	00F	PL B484 198	D. Barberis et al.	(WA 102 Collab.)
ADOMEIT	96	ZPHY C71 227	J. Adomeit et al.	(Crystal Barrel Collab.)
BELADIDZE	92B	ZPHY C54 367	G.M. Beladidze et al.	(VES Collab.)
BELADIDZE	92D	ZPHY C57 13	G.M. Beladidze et al.	(VES Collab.)
ALDE	91B	SJNP 54 455	D.M. Alde et al.	(SERP, BELG, LANL, LAPP+)
		Translated from YAF 54 751.		
Also		PL B276 375	D.M. Alde et al.	(BELG, SERP, KEK, LANL+)
ALDE	90	PL B241 600	D.M. Alde et al.	(SERP, BELG, LANL, LAPP+)
ALDE	89	PL B216 447	D.M. Alde et al.	(SERP, BELG, LANL, LAPP+)
Also		SJNP 48 1035	D.M. Alde et al.	(BELG, SERP, LANL, LAPP)
		Translated from YAF 48 1724.		
BALOSHIN	86	SJNP 43 959	O.N. Baloshin et al.	(ITEP)
		Translated from YAF 43 1487.		

### $a_0(1950)$

$$I^G(J^{PC}) = 1^-(0^{++})$$

OMITTED FROM SUMMARY TABLE

Needs confirmation. Seen in  $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$  by LEES 16A with significance 2.5  $\sigma$  in  $K_S^0 K^\pm \pi^\mp$  and 4.2  $\sigma$  in  $K^+ K^- \pi^0$ . Spin-2 explanation ( $a_2(1950)$ ) is not compatible with data.

### $a_0(1950)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1931±14±22</b>	12k	<sup>1,2</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1949±32±76	8k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K_S^0 K^\pm \pi^\mp$
1927±15±23	4k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K^+ K^- \pi^0$
<sup>1</sup> From a model-independent partial wave analysis fit to a relativistic Breit-Wigner function with a floating width.				
<sup>2</sup> Weighted average of the $K_S^0 K^\pm$ and $K^+ K^-$ decay modes.				

### $a_0(1950)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>271±22±29</b>	12k	<sup>1,2</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
265±36±110	8k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K_S^0 K^\pm \pi^\mp$
274±28±30	4k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K^+ K^- \pi^0$
<sup>1</sup> From a model-independent partial wave analysis fit to a relativistic Breit-Wigner function with a floating mass.				
<sup>2</sup> Weighted average of the $K_S^0 K^\pm$ and $K^+ K^-$ decay modes.				

# Meson Particle Listings

## $a_0(1950)$ , $f_2(1950)$

$a_0(1950)$ DECAY MODES				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$K\bar{K}$	seen		

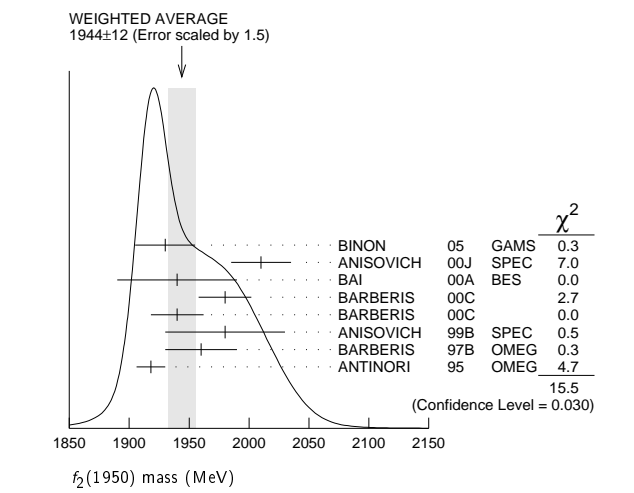
$a_0(1950)$ BRANCHING RATIOS				
$\Gamma(K\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
seen	12k	<sup>1</sup> LEES	16A BABR	$\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\bar{K}\pi$

<sup>1</sup> From a model-independent partial wave analysis.

$a_0(1950)$ REFERENCES				
LEES	16A	PR D93 012005	J.P. Lees <i>et al.</i>	(BABAR Collab.)

$f_2(1950)$	$I^G(J^{PC}) = 0^+(2^{++})$
-------------	-----------------------------

$f_2(1950)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1944±12 OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.			
1930±25	<sup>1</sup> BINON	05	GAMS	33 $\pi^-p \rightarrow \eta\eta n$
2010±25	ANISOVICH	00J	SPEC	
1940±50	BAI	00A	BES	$J/\psi \rightarrow \gamma(\pi^+\pi^-\pi^+\pi^-)$
1980±22	<sup>2</sup> BARBERIS	00C	450 $pp \rightarrow pp4\pi$	
1940±22	<sup>3</sup> BARBERIS	00C	450 $pp \rightarrow pp2\pi2\pi^0$	
1980±50	ANISOVICH	99B	SPEC	1.35-1.94 $p\bar{p} \rightarrow \eta\eta\pi^0$
1960±30	BARBERIS	97B	OMEG	450 $pp \rightarrow pp2(\pi^+\pi^-)$
1918±12	ANTINORI	95	OMEG	300,450 $pp \rightarrow pp2(\pi^+\pi^-)$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
2038 <sup>+13+12</sup> -11-73	<sup>4</sup> UEHARA	09	BELL	10.6 $e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
1980± 2±14	ABE	04	BELL	10.6 $e^+e^- \rightarrow e^+e^-K^+K^-$
1867±46	<sup>5</sup> AMSLER	02	CBAR	0.9 $\bar{p}p \rightarrow \pi^0\eta\eta, \pi^0\pi^0\pi^0$
~ 1990	<sup>6</sup> OAKDEN	94	RVUE	0.36-1.55 $\bar{p}p \rightarrow \pi\pi$
1950±15	<sup>7</sup> ASTON	91	LASS	11 $K^-p \rightarrow \Lambda K\bar{K}\pi\pi$
<sup>1</sup> First solution, PWA is ambiguous. <sup>2</sup> Decaying into $\pi^+\pi^-\pi^0$ . <sup>3</sup> Decaying into $2(\pi^+\pi^-)$ . <sup>4</sup> Taking into account $f_4(2050)$ . <sup>5</sup> T-matrix pole. <sup>6</sup> From solution B of amplitude analysis of data on $\bar{p}p \rightarrow \pi\pi$ . See however KLOET 96 who fit $\pi^+\pi^-$ only and find waves only up to $J=3$ to be important but not significantly resonant. <sup>7</sup> Cannot determine spin to be 2.				



$f_2(1950)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>472± 18 OUR AVERAGE</b>				
450± 50	<sup>8</sup> BINON	05	GAMS	33 $\pi^- p \rightarrow \eta \eta n$
495± 35	ANISOVICH	00J	SPEC	
380 <sup>+120</sup> <sub>- 90</sub>	BAI	00A	BES	$J/\psi \rightarrow \gamma(\pi^+ \pi^- \pi^+ \pi^-)$
520± 50	<sup>9</sup> BARBERIS	00C		450 $pp \rightarrow pp4\pi$
485± 55	<sup>10</sup> BARBERIS	00C		450 $pp \rightarrow pp4\pi$
500±100	ANISOVICH	99B	SPEC	1.35-1.94 $p\bar{p} \rightarrow \eta \eta \pi^0$
460± 40	BARBERIS	97B	OMEG	450 $pp \rightarrow pp2(\pi^+ \pi^-)$
390± 60	ANTINORI	95	OMEG	300,450 $pp \rightarrow pp2(\pi^+ \pi^-)$

• • • We do not use the following data for averages, fits, limits, etc. • • •				
441 <sup>+27+</sup> <sub>-25-192</sub>	<sup>11</sup> UEHARA	09	BELL	$10.6\ e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
297±12±6	ABE	04	BELL	$10.6\ e^+e^- \rightarrow e^+e^-K^+K^-$
385±58	<sup>12</sup> AMSLER	02	CBAR	$0.9\ \bar{p}p \rightarrow \pi^0\eta\eta, \pi^0\pi^0\pi^0$
~100	<sup>13</sup> OAKDEN	94	RVUE	$0.36\text{--}1.55\ \bar{p}p \rightarrow \pi\pi$
250±50	<sup>14</sup> ASTON	91	LASS	$11\ K^-\rho \rightarrow \Lambda K\bar{K}\pi\pi$
<sup>8</sup> First solution, PWA is ambiguous. <sup>9</sup> Decaying into $\pi^+\pi^-2\pi^0$ . <sup>10</sup> Decaying into $2(\pi^+\pi^-)$ . <sup>11</sup> Taking into account $f_4(2050)$ . <sup>12</sup> T-matrix pole. <sup>13</sup> From solution B of amplitude analysis of data on $\bar{p}p \rightarrow \pi\pi$ . See however KLOET 96 who fit $\pi^+\pi^-$ only and find waves only up to $J=3$ to be important but not significantly resonant. <sup>14</sup> Cannot determine spin to be 2.				

$f_2(1950)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $K^*(892)\bar{K}^*(892)$	seen	
$\Gamma_2$ $\pi\pi$		
$\Gamma_3$ $\pi^+\pi^-$	seen	
$\Gamma_4$ $\pi^0\pi^0$	seen	
$\Gamma_5$ $4\pi$	seen	
$\Gamma_6$ $\pi^+\pi^-\pi^+\pi^-$		
$\Gamma_7$ $a_2(1320)\pi$		
$\Gamma_8$ $f_2(1270)\pi\pi$		
$\Gamma_9$ $\eta\eta$	seen	
$\Gamma_{10}$ $K\bar{K}$	seen	
$\Gamma_{11}$ $\gamma\gamma$	seen	
$\Gamma_{12}$ $\rho\bar{\rho}$	seen	

$f_2(1950)$ $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$		
$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_{10}\Gamma_{11}/\Gamma$	
VALUE (eV)	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
122±4±26	<sup>15</sup> ABE	04 BELL $10.6\ e^+e^- \rightarrow e^+e^-K^+K^-$
<sup>15</sup> Assuming spin 2.		

$\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$		
$\Gamma_2\Gamma_{11}/\Gamma$		
VALUE	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
162 <sup>+69+1137</sup> <sub>-42-204</sub>	<sup>16</sup> UEHARA	09 BELL $10.6\ e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
<sup>16</sup> Taking into account $f_4(2050)$ .		

$f_2(1950)$ BRANCHING RATIOS				
$\Gamma(K^*(892)\bar{K}^*(892))/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	ASTON	91	LASS	0 11 $K^-\rho \rightarrow \Lambda K\bar{K}\pi\pi$
$\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen	BARBERIS	00B	450	$p\rho \rightarrow p_f\eta\pi^+\pi^-\rho_s$
not seen	BARBERIS	00C	450	$p\rho \rightarrow p_f4\pi\rho_s$
possibly seen	BARBERIS	97B	OMEG	450 $p\rho \rightarrow p\rho2(\pi^+\pi^-)$
$\Gamma(\eta\eta)/\Gamma(4\pi)$				$\Gamma_9/\Gamma_5$
VALUE	CL%	DOCUMENT ID	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5.0 \times 10^{-3}$	90	BARBERIS	00E	450 $p\rho \rightarrow p_f\eta\eta\rho_s$
$\Gamma(\eta\eta)/\Gamma(\pi^+\pi^-)$				$\Gamma_9/\Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.14 \pm 0.05</math></b>	AMSLER	02	CBAR	$0.9\ \bar{p}p \rightarrow \pi^0\eta\eta, \pi^0\pi^0\pi^0$
$\Gamma(p\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
seen	111	ALEXANDER	10	CLEO $\psi(2S) \rightarrow \gamma p\bar{p}$



See key on page 601

## Meson Particle Listings

 $f_2(1950)$ ,  $\rho_3(1990)$ ,  $f_2(2010)$ ,  $f_0(2020)$  $f_2(1950)$  REFERENCES

ALEXANDER	10	PR D82 092002	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
UEHARA	09	PR D79 052009	S. Uehara <i>et al.</i>	(BELLE Collab.)
BINON	05	PAN 68 960	F. Binon <i>et al.</i>	
		Translated from YAF 68 998.		
ABE	04	EPJ C32 323	K. Abe <i>et al.</i>	(BELLE Collab.)
AMSLER	02	EPJ C23 29	C. Amisler <i>et al.</i>	
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
BAI	00A	PL B472 207	J.Z. Bai <i>et al.</i>	(BES Collab.)
BARBERIS	00B	PL B471 435	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00C	PL B471 440	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS	00E	PL B479 59	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ANISOVICH	99B	PL B449 154	A.V. Anisovich <i>et al.</i>	
BARBERIS	97B	PL B413 217	D. Barberis <i>et al.</i>	(WA 102 Collab.)
KLOET	96	PR D53 6120	W.M. Kloet, F. Myhrer	(RUTG, NORD)
ANTINORI	95	PL B353 589	F. Antinori <i>et al.</i>	(ATHU, BARI, BIRM+)JP
OAKDEN	94	NP A574 731	M.N. Oakden, M.R. Pennington	(DURH)
ASTON	91	NPBPS B21 5	D. Aston <i>et al.</i>	(LASS Collab.)

 $\rho_3(1990)$ 

$$J^G(J^{PC}) = 1^+(3^{--})$$

OMITTED FROM SUMMARY TABLE

 $\rho_3(1990)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1982 \pm 14$	<sup>1</sup> ANISOVICH 02	SPEC	$0.6\text{--}1.9 \, p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$
~ 2007	HASAN 94	RVUE	$\bar{p}p \rightarrow \pi\pi$
<sup>1</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.			

 $\rho_3(1990)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$188 \pm 24$	<sup>2</sup> ANISOVICH 02	SPEC	$0.6\text{--}1.9 \, p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$
~ 287	HASAN 94	RVUE	$\bar{p}p \rightarrow \pi\pi$
<sup>2</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.			

 $\rho_3(1990)$  REFERENCES

ANISOVICH	02	PL B542 8	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01D	PL B508 6	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01E	PL B513 281	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
HASAN	94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)

 $f_2(2010)$ 

$$J^G(J^{PC}) = 0^+(2^{++})$$

 $f_2(2010)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$2011 \pm \frac{62}{76}$	<sup>1</sup> ETKIN 88	MPS	$22 \, \pi^-\pi^- \rightarrow \phi\phi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2005 \pm 12$	VLADIMIRSK...06	SPEC	$40 \, \pi^-\pi^- \rightarrow K_S^0 K_S^0 n$
$1980 \pm 20$	<sup>2</sup> BOLONKIN 88	SPEC	$40 \, \pi^-\pi^- \rightarrow K_S^0 K_S^0 n$
$2050 \pm \frac{90}{50}$	ETKIN 85	MPS	$22 \, \pi^-\pi^- \rightarrow 2\phi n$
$2120 \pm \frac{20}{120}$	LINDENBAUM 84	RVUE	
$2160 \pm 50$	ETKIN 82	MPS	$22 \, \pi^-\pi^- \rightarrow 2\phi n$
<sup>1</sup> Includes data of ETKIN 85. The percentage of the resonance going into $\phi\phi \, 2^++ \, S_2$ , $D_2$ , and $D_0$ is $98 \pm \frac{1}{3}$ , $0 \pm \frac{1}{0}$ , and $2 \pm \frac{2}{1}$ , respectively.			
<sup>2</sup> Statistically very weak, only 1.4 s.d.			

 $f_2(2010)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$202 \pm \frac{67}{62}$	<sup>3</sup> ETKIN 88	MPS	$22 \, \pi^-\pi^- \rightarrow \phi\phi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$209 \pm 32$	VLADIMIRSK...06	SPEC	$40 \, \pi^-\pi^- \rightarrow K_S^0 K_S^0 n$
$145 \pm 50$	<sup>4</sup> BOLONKIN 88	SPEC	$40 \, \pi^-\pi^- \rightarrow K_S^0 K_S^0 n$
$200 \pm \frac{160}{50}$	ETKIN 85	MPS	$22 \, \pi^-\pi^- \rightarrow 2\phi n$
$300 \pm \frac{150}{50}$	LINDENBAUM 84	RVUE	
$310 \pm 70$	ETKIN 82	MPS	$22 \, \pi^-\pi^- \rightarrow 2\phi n$
<sup>3</sup> Includes data of ETKIN 85.			
<sup>4</sup> Statistically very weak, only 1.4 s.d.			

 $f_2(2010)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \, \phi\phi$	seen
$\Gamma_2 \, K\bar{K}$	seen

 $f_2(2010)$  BRANCHING RATIOS

$\Gamma(K\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
seen	VLADIMIRSK...06	SPEC	$40 \, \pi^-\pi^- \rightarrow K_S^0 K_S^0 n$	

 $f_2(2010)$  REFERENCES

VLADIMIRSK... 06	PAN 69 493	V.V. Vladimirov <i>et al.</i>	(ITEP, Moscow)
	Translated from YAF 69 515.		
BOLONKIN 88	NP B309 426	B.V. Bolonkin <i>et al.</i>	(ITEP, SERP)
ETKIN 88	PL B201 568	A. Etkin <i>et al.</i>	(BNL, CUNY)
ETKIN 85	PL B165B 217	A. Etkin <i>et al.</i>	(BNL, CUNY)
LINDENBAUM 84	CNPP 13 285	S.J. Lindenbaum	(CUNY)
ETKIN 82	PRL 49 1620	A. Etkin <i>et al.</i>	(BNL, CUNY)
Also	Brighton Conf. 351	S.J. Lindenbaum	(BNL, CUNY)

 $f_0(2020)$ 

$$J^G(J^{PC}) = 0^+(0^{++})$$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

 $f_0(2020)$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1992 ± 16</b>	<sup>1,2</sup> BARBERIS 00c			$450 \, p\bar{p} \rightarrow p_f 4\pi p_S$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2037 \pm 8$	80k	<sup>3</sup> UMAN 06	E835	$5.2 \, \bar{p}p \rightarrow \eta\eta\pi^0$
$2040 \pm 38$		ANISOVICH 00J	SPEC	
$2010 \pm 60$		ALDE 98	GAM4	$100 \, \pi^-\pi^- \rightarrow \pi^0\pi^0 n$
$2020 \pm 35$		BARBERIS 97B	OMEG	$450 \, p\bar{p} \rightarrow p\rho 2(\pi^+\pi^-)$
<sup>1</sup> Average between $\pi^+\pi^- 2\pi^0$ and $2(\pi^+\pi^-)$ .				
<sup>2</sup> T-matrix pole.				
<sup>3</sup> Statistical error only.				

 $f_0(2020)$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>442 ± 60</b>	<sup>4,5</sup> BARBERIS 00c			$450 \, p\bar{p} \rightarrow p_f 4\pi p_S$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$296 \pm 17$	80k	<sup>6</sup> UMAN 06	E835	$5.2 \, \bar{p}p \rightarrow \eta\eta\pi^0$
$405 \pm 40$		ANISOVICH 00J	SPEC	
$240 \pm 100$		ALDE 98	GAM4	$100 \, \pi^-\pi^- \rightarrow \pi^0\pi^0 n$
$410 \pm 50$		BARBERIS 97B	OMEG	$450 \, p\bar{p} \rightarrow p\rho 2(\pi^+\pi^-)$
<sup>4</sup> Average between $\pi^+\pi^- 2\pi^0$ and $2(\pi^+\pi^-)$ .				
<sup>5</sup> T-matrix pole.				
<sup>6</sup> Statistical error only.				

 $f_0(2020)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \, \rho\pi\pi$	seen
$\Gamma_2 \, \pi^0\pi^0$	seen
$\Gamma_3 \, \rho\rho$	seen
$\Gamma_4 \, \omega\omega$	seen
$\Gamma_5 \, \eta\eta$	seen

 $f_0(2020)$  BRANCHING RATIOS

$\Gamma(\rho\rho)/\Gamma(\omega\omega)$				$\Gamma_3/\Gamma_4$
VALUE	DOCUMENT ID	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 3	BARBERIS 00F	450	$p\bar{p} \rightarrow p_f \omega \omega p_S$	
$\Gamma(\eta\eta)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	UMAN 06	E835	$5.2 \, p\bar{p} \rightarrow \eta\eta\pi^0$	

 $f_0(2020)$  REFERENCES

UMAN 06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
ANISOVICH 00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
BARBERIS 00C	PL B471 440	D. Barberis <i>et al.</i>	(WA 102 Collab.)
BARBERIS 00F	PL B484 198	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ALDE 98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
Also	PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
	Translated from YAF 62 446.		
BARBERIS 97B	PL B413 217	D. Barberis <i>et al.</i>	(WA 102 Collab.)

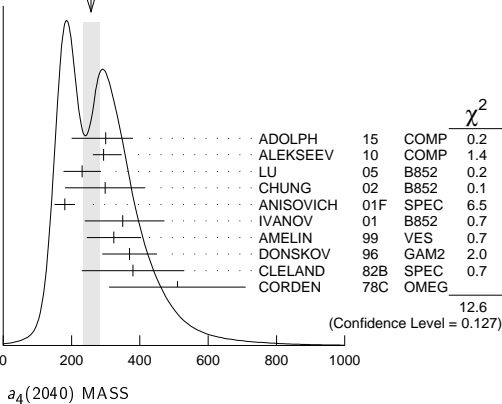
Meson Particle Listings

$a_4(2040)$ ,  $f_4(2050)$

$a_4(2040)$					
$I^G(J^{PC}) = 1^-(4^{++})$					
$a_4(2040)$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1995<math>\pm</math>10<math>\pm</math>8 OUR AVERAGE</b>		Error includes scale factor of 1.1.			
1900 $\pm$ 80 $\pm$ 20		ADOLPH	15	COMP	191 $\pi^- p \rightarrow \eta^{(\prime)} \pi^- p$
1885 $\pm$ 13 $\pm$ 50 $\pm$ 2	420k	ALEKSEEV	10	COMP	190 $\pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
1985 $\pm$ 10 $\pm$ 13	145k	LU	05	B852	18 $\pi^- p \rightarrow \omega \pi^- \pi^0 p$
1996 $\pm$ 25 $\pm$ 43		CHUNG	02	B852	18.3 $\pi^- p \rightarrow 3\pi p$
2005 $\pm$ 25 $\pm$ 45		<sup>1</sup> ANISOVICH	01F	SPEC	2.0 $\overline{p} p \rightarrow 3\pi^0, \pi^0 \eta, \pi^0 \eta'$
2000 $\pm$ 40 $\pm$ 60 $\pm$ 20		IVANOV	01	B852	18 $\pi^- p \rightarrow \eta' \pi^- p$
1944 $\pm$ 8 $\pm$ 50		<sup>2</sup> AMELIN	99	VES	37 $\pi^- A \rightarrow \omega \pi^- \pi^0 A^*$
2010 $\pm$ 20		<sup>3</sup> DONSKOV	96	GAM2 0	38 $\pi^- p \rightarrow \eta \pi^0 n$
2040 $\pm$ 30		<sup>4</sup> CLELAND	82b	SPEC $\pm$	50 $\pi p \rightarrow K_S^0 K^\pm p$
2030 $\pm$ 50		<sup>5</sup> CORDEN	78c	OMEG 0	15 $\pi^- p \rightarrow 3\pi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2004 $\pm$ 6	80k	<sup>6</sup> UMAN	06	E835	5.2 $\overline{p} p \rightarrow \eta \eta \pi^0$
1903 $\pm$ 10		<sup>7</sup> BALDI	78	SPEC $-$	10 $\pi^- p \rightarrow \rho K_S^0 K^-$
<sup>1</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, and ANISOVICH 01F. <sup>2</sup> May be a different state. <sup>3</sup> From a simultaneous fit to the $G_+$ and $G_0$ wave intensities. <sup>4</sup> From an amplitude analysis. <sup>5</sup> $J^P = 4^+$ is favored, though $J^P = 2^+$ cannot be excluded. <sup>6</sup> Statistical error only. <sup>7</sup> From a fit to the $Y_8^0$ moment. Limited by phase space.					

$a_4(2040)$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>257<math>\pm</math>25<math>\pm</math>23 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.			
300 $\pm$ 80 $\pm$ 100		ADOLPH	15	COMP	191 $\pi^- p \rightarrow \eta^{(\prime)} \pi^- p$
294 $\pm$ 25 $\pm$ 46 $\pm$ 19	420k	ALEKSEEV	10	COMP	190 $\pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
231 $\pm$ 30 $\pm$ 46	145k	LU	05	B852	18 $\pi^- p \rightarrow \omega \pi^- \pi^0 p$
298 $\pm$ 81 $\pm$ 85		CHUNG	02	B852	18.3 $\pi^- p \rightarrow 3\pi p$
180 $\pm$ 30		<sup>1</sup> ANISOVICH	01F	SPEC	2.0 $\overline{p} p \rightarrow 3\pi^0, \pi^0 \eta, \pi^0 \eta'$
350 $\pm$ 100 $\pm$ 70 $\pm$ 50		IVANOV	01	B852	18 $\pi^- p \rightarrow \eta' \pi^- p$
324 $\pm$ 26 $\pm$ 75		<sup>2</sup> AMELIN	99	VES	37 $\pi^- A \rightarrow \omega \pi^- \pi^0 A^*$
370 $\pm$ 80		<sup>3</sup> DONSKOV	96	GAM2 0	38 $\pi^- p \rightarrow \eta \pi^0 n$
380 $\pm$ 150		<sup>4</sup> CLELAND	82b	SPEC $\pm$	50 $\pi p \rightarrow K_S^0 K^\pm p$
510 $\pm$ 200		<sup>5</sup> CORDEN	78c	OMEG 0	15 $\pi^- p \rightarrow 3\pi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
401 $\pm$ 16	80k	<sup>6</sup> UMAN	06	E835	5.2 $\overline{p} p \rightarrow \eta \eta \pi^0$
166 $\pm$ 43		<sup>7</sup> BALDI	78	SPEC $-$	10 $\pi^- p \rightarrow \rho K_S^0 K^-$
<sup>1</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, and ANISOVICH 01F. <sup>2</sup> May be a different state. <sup>3</sup> From a simultaneous fit to the $G_+$ and $G_0$ wave intensities. <sup>4</sup> From an amplitude analysis. <sup>5</sup> $J^P = 4^+$ is favored, though $J^P = 2^+$ cannot be excluded. <sup>6</sup> Statistical error only. <sup>7</sup> From a fit to the $Y_8^0$ moment. Limited by phase space.					

WEIGHTED AVERAGE  
257 $\pm$ 25-23 (Error scaled by 1.3)



$a_4(2040)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\overline{K}$	seen
$\Gamma_2$ $\pi^+ \pi^- \pi^0$	seen
$\Gamma_3$ $\rho \pi$	seen
$\Gamma_4$ $f_2(1270) \pi$	seen
$\Gamma_5$ $\omega \pi^- \pi^0$	seen
$\Gamma_6$ $\omega \rho$	seen
$\Gamma_7$ $\eta \pi$	seen
$\Gamma_8$ $\eta'(958) \pi$	seen

$a_4(2040)$ BRANCHING RATIOS				
$\Gamma(K\overline{K})/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	BALDI	78	SPEC $\pm$	10 $\pi^- p \rightarrow K_S^0 K^- p$
$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	CORDEN	78c	OMEG 0	15 $\pi^- p \rightarrow 3\pi n$
$\Gamma(\rho \pi)/\Gamma(f_2(1270) \pi)$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
1.1 $\pm$ 0.2 $\pm$ 0.2	CHUNG	02	B852	18.3 $\pi^- p \rightarrow 3\pi p$
$\Gamma(\eta \pi)/\Gamma_{\text{total}}$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	DONSKOV	96	GAM2 0	38 $\pi^- p \rightarrow \eta \pi^0 n$

$\Gamma(\eta'(958) \pi)/\Gamma(\eta \pi)$			
VALUE	DOCUMENT ID	TECN	COMMENT
0.23 $\pm$ 0.07	ADOLPH	15	COMP 191 $\pi^- p \rightarrow \eta^{(\prime)} \pi^- p$
$\Gamma(\omega \rho)/\Gamma_{\text{total}}$			
VALUE	EVTS	DOCUMENT ID	TECN
seen	145k	LU	05

$a_4(2040)$ REFERENCES			
ADOLPH	15	PL B740 303	M. Adolph <i>et al.</i> (COMPASS Collab.)
ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i> (COMPASS Collab.)
UMAN	06	PR D73 052009	I. Uman <i>et al.</i> (FNAL E835)
LU	05	PRL 94 032002	M. Lu <i>et al.</i> (BNL E852 Collab.)
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i> (BNL E852 Collab.)
ANISOVICH	01F	PL B517 261	A.V. Anisovich <i>et al.</i>
IVANOV	01	PRL 86 3977	E.I. Ivanov <i>et al.</i> (BNL E852 Collab.)
AMELIN	99	PAN 62 445	D.V. Amelin <i>et al.</i> (VES Collab.)
ANISOVICH	99C	Translated from YAF 62 487.	A.V. Anisovich <i>et al.</i>
ANISOVICH	99E	PL B452 173	A.V. Anisovich <i>et al.</i>
DONSKOV	96	PAN 59 982	S.V. Donskov <i>et al.</i> (GAMS Collab.)
CLELAND	82B	NP B208 228	W.E. Cleland <i>et al.</i> (DURH, GEVA, LAUS+)
BALDI	78	PL 74B 413	R. Baldi <i>et al.</i> (GEVA)JP
CORDEN	78C	NP B136 77	M.J. Corden <i>et al.</i> (BIRM, RHEL, TELA+)JP

$f_4(2050)$					
$I^G(J^{PC}) = 0^+(4^{++})$					
$f_4(2050)$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2018<math>\pm</math>11 OUR AVERAGE</b>		Error includes scale factor of 2.1. See the ideogram below.			
1960 $\pm$ 15		AMELIN	06	VES	36 $\pi^- p \rightarrow \omega \omega n$
2005 $\pm$ 10		<sup>1</sup> BINON	05	GAMS	33 $\pi^- p \rightarrow \eta \eta n$
1998 $\pm$ 15		ALDE	98	GAM4	100 $\pi^- p \rightarrow \pi^0 \pi^0 n$
2060 $\pm$ 20		ALDE	90	GAM2	38 $\pi^- p \rightarrow \omega \omega n$
2038 $\pm$ 30		AUGUSTIN	87	DM2	$J/\psi \rightarrow \gamma \pi^+ \pi^-$
2086 $\pm$ 15		BALTRUSAIT..	87	MRK3	$J/\psi \rightarrow \gamma \pi^+ \pi^-$
2000 $\pm$ 60		ALDE	86D	GAM4	100 $\pi^- p \rightarrow n 2\eta$
2020 $\pm$ 20	40k	<sup>2</sup> BINON	84B	GAM2	38 $\pi^- p \rightarrow n 2\pi^0$
2015 $\pm$ 28		<sup>3</sup> CASON	82	STRC	8 $\pi^+ p \rightarrow \Delta^+ \pi^0 \pi^0$
2031 $\pm$ 25 $\pm$ 36		ETKIN	82B	MPS	23 $\pi^- p \rightarrow n 2K_S^0$
2020 $\pm$ 30	700	APEL	75	NICE	40 $\pi^- p \rightarrow n 2\pi^0$
2050 $\pm$ 25		BLUM	75	ASPK	18.4 $\pi^- p \rightarrow n K^+ K^-$

See key on page 601

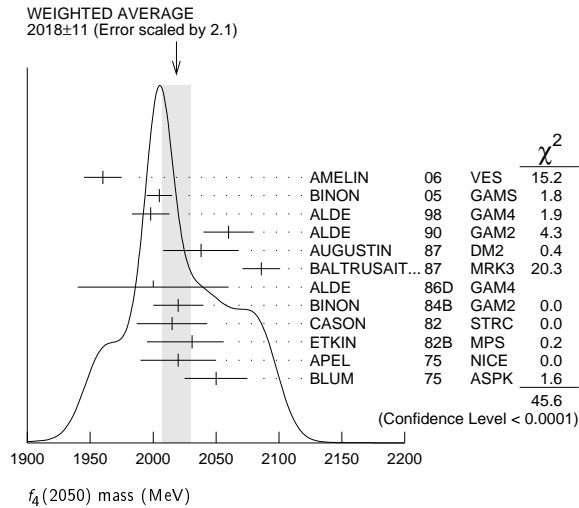
## Meson Particle Listings

 $f_4(2050)$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

1966±25	4	ANISOVICH	09	RVUE	0.0 $\bar{p}p, \pi N$
1885 <sup>+14+218</sup> <sub>-13-25</sub>	5	UEHARA	09	BELL	10.6 $e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
2018±6		ANISOVICH	00J	SPEC	2.0 $\bar{p}p \rightarrow \eta\pi^0\pi^0, \pi^0\pi^0,$ $\eta\eta, \eta\eta', \pi\pi$
~2000	6	MARTIN	98	RVUE	$N\bar{N} \rightarrow \pi\pi$
~2010	7	MARTIN	97	RVUE	$N\bar{N} \rightarrow \pi\pi$
~2040	8	OAKDEN	94	RVUE	0.36-1.55 $\bar{p}p \rightarrow \pi\pi$
~1990	9	OAKDEN	94	RVUE	0.36-1.55 $\bar{p}p \rightarrow \pi\pi$
1978±5	10	ALPER	80	CNTR	62 $\pi^-p \rightarrow K^+K^-\pi$
2040±10	10	ROZANSKA	80	SPRK	18 $\pi^-p \rightarrow p\bar{p}n$
1935±13	10	CORDEN	79	OMEG	12-15 $\pi^-p \rightarrow n2\pi$
1988±7		EVANGELIS...	79B	OMEG	10 $\pi^-p \rightarrow K^+K^-\pi$
1922±14	11	ANTIPOV	77	CIBS	25 $\pi^-p \rightarrow p3\pi$

- 1 From the first PWA solution.  
2 From a partial-wave analysis of the data.  
3 From an amplitude analysis of the reaction  $\pi^+\pi^- \rightarrow 2\pi^0$ .  
4 K matrix pole.  
5 Taking into account the  $f_2(1950)$ . Helicity-2 production favored.  
6 Energy-dependent analysis.  
7 Single energy analysis.  
8 From solution A of amplitude analysis of data on  $\bar{p}p \rightarrow \pi\pi$ . See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J=3$  to be important but not significantly resonant.  
9 From solution B of amplitude analysis of data on  $\bar{p}p \rightarrow \pi\pi$ . See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J=3$  to be important but not significantly resonant.  
10  $I(J^P) = 0(4^+)$  from amplitude analysis assuming one-pion exchange.  
11 Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

 $f_4(2050)$  WIDTH

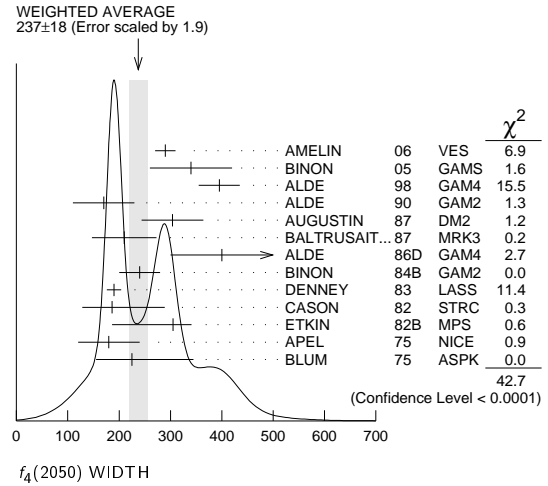
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>237±18 OUR AVERAGE</b>		Error includes scale factor of 1.9. See the ideogram below.		
290±20		AMELIN	06	VES 36 $\pi^-p \rightarrow \omega\omega n$
340±80	12	BINON	05	GAMS 33 $\pi^-p \rightarrow \eta\eta n$
395±40		ALDE	98	GAM4 100 $\pi^-p \rightarrow \pi^0\pi^0 n$
170±60		ALDE	90	GAM2 38 $\pi^-p \rightarrow \omega\omega n$
304±60		AUGUSTIN	87	DM2 $J/\psi \rightarrow \gamma\pi^+\pi^-$
210±63		BALTRUSAIT...	87	MRK3 $J/\psi \rightarrow \gamma\pi^+\pi^-$
400±100		ALDE	86D	GAM4 100 $\pi^-p \rightarrow n2\eta$
240±40	13	BINON	84B	GAM2 38 $\pi^-p \rightarrow n2\pi^0$
190±14		DENNEY	83	LASS 10 $\pi^+n/\pi^+p$
186 <sup>+103</sup> <sub>-58</sub>	14	CASON	82	STRC 8 $\pi^+p \rightarrow \Delta^{++}\pi^0\pi^0$
305 <sup>+36</sup> <sub>-119</sub>		ETKIN	82B	MPS 23 $\pi^-p \rightarrow n2K_S^0$
180±60	700	APEL	75	NICE 40 $\pi^-p \rightarrow n2\pi^0$
225 <sup>+120</sup> <sub>-70</sub>		BLUM	75	ASPK 18.4 $\pi^-p \rightarrow nK^+K^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

260±40	15	ANISOVICH	09	RVUE	0.0 $\bar{p}p, \pi N$
453±20 <sup>+31</sup> <sub>-129</sub>	16	UEHARA	09	BELL	10.6 $e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
182±7		ANISOVICH	00J	SPEC	2.0 $\bar{p}p \rightarrow \eta\pi^0\pi^0, \pi^0\pi^0,$ $\eta\eta, \eta\eta', \pi\pi$
~170	17	MARTIN	98	RVUE	$N\bar{N} \rightarrow \pi\pi$
~200	18	MARTIN	97	RVUE	$N\bar{N} \rightarrow \pi\pi$
~60	19	OAKDEN	94	RVUE	0.36-1.55 $\bar{p}p \rightarrow \pi\pi$

~80	20	OAKDEN	94	RVUE	0.36-1.55 $\bar{p}p \rightarrow \pi\pi$
243±16	21	ALPER	80	CNTR	62 $\pi^-p \rightarrow K^+K^-\pi$
140±15	21	ROZANSKA	80	SPRK	18 $\pi^-p \rightarrow p\bar{p}n$
263±57	21	CORDEN	79	OMEG	12-15 $\pi^-p \rightarrow n2\pi$
100±28		EVANGELIS...	79B	OMEG	10 $\pi^-p \rightarrow K^+K^-\pi$
107±56	22	ANTIPOV	77	CIBS	25 $\pi^-p \rightarrow p3\pi$

- 12 From the first PWA solution.  
13 From a partial-wave analysis of the data.  
14 From an amplitude analysis of the reaction  $\pi^+\pi^- \rightarrow 2\pi^0$ .  
15 K matrix pole.  
16 Taking into account the  $f_2(1950)$ . Helicity-2 production favored.  
17 Energy-dependent analysis.  
18 Single energy analysis.  
19 From solution A of amplitude analysis of data on  $\bar{p}p \rightarrow \pi\pi$ . See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J=3$  to be important but not significantly resonant.  
20 From solution B of amplitude analysis of data on  $\bar{p}p \rightarrow \pi\pi$ . See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J=3$  to be important but not significantly resonant.  
21  $I(J^P) = 0(4^+)$  from amplitude analysis assuming one-pion exchange.  
22 Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

 $f_4(2050)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\omega\omega$	seen
$\Gamma_2$ $\pi\pi$	(17.0±1.5) %
$\Gamma_3$ $K\bar{K}$	( 6.8 <sup>+3.4</sup> <sub>-1.8</sub> ) × 10 <sup>-3</sup>
$\Gamma_4$ $\eta\eta$	( 2.1±0.8) × 10 <sup>-3</sup>
$\Gamma_5$ $4\pi^0$	< 1.2 %
$\Gamma_6$ $\gamma\gamma$	
$\Gamma_7$ $a_2(1320)\pi$	seen

 $f_4(2050)$   $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_3\Gamma_6/\Gamma$
--	---------------------------

VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.29	95	ALTHOFF	85B	TASS $\gamma\gamma \rightarrow K\bar{K}\pi$

$\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_2\Gamma_6/\Gamma$
--	---------------------------

VALUE (eV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
23.1 <sup>+3.6+70.5</sup> <sub>-3.3-15.6</sub>			23 UEHARA	09	BELL 10.6 $e^+e^- \rightarrow e^+e^-\pi^0\pi^0$
<1100	95	13±4	OEST	90	JADE $e^+e^- \rightarrow e^+e^-\pi^0\pi^0$

23 Taking into account the  $f_2(1950)$ . Helicity-2 production favored. $f_4(2050)$  BRANCHING RATIOS

$\Gamma(\omega\omega)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
seen	AMELIN 06 VES 36 $\pi^-p \rightarrow \omega\omega n$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
not seen	BARBERIS 00F 450 $pp \rightarrow p f \omega p_S$

# Meson Particle Listings

## $f_4(2050)$ , $\pi_2(2100)$ , $f_0(2100)$

$\Gamma(\omega\omega)/\Gamma(\pi\pi)$	$\Gamma_1/\Gamma_2$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.5\pm0.3</math></b>	ALDE	90	GAM2	$38\pi^-p\rightarrow\omega\omega n$
$\Gamma(\pi\pi)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.170\pm0.015</math> OUR AVERAGE</b>				
$0.18\pm0.03$	24 BINON	83c	GAM2	$38\pi^-p\rightarrow n4\gamma$
$0.16\pm0.03$	24 CASON	82	STRC	$8\pi^+p\rightarrow\Delta^{++}\pi^0\pi^0$
$0.17\pm0.02$	24 CORDEN	79	OMEG	$12\text{--}15\pi^-p\rightarrow n2\pi$
24 Assuming one pion exchange.				
$\Gamma(K\overline{K})/\Gamma(\pi\pi)$	$\Gamma_3/\Gamma_2$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.04^{+0.02}_{-0.01}</math></b>	ETKIN	82b	MPS	$23\pi^-p\rightarrow n2K_S^0$
$\Gamma(\eta\eta)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$			
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>2.1\pm0.8</math></b>	ALDE	86d	GAM4	$100\pi^-p\rightarrow n4\gamma$
$\Gamma(4\pi^0)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;0.012</math></b>	ALDE	87	GAM4	$100\pi^-p\rightarrow 4\pi^0 n$
$\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>seen</b>	AMELIN	00	VES	$37\pi^-p\rightarrow\eta\pi^+\pi^-n$

### $f_4(2050)$ REFERENCES

ANISOVICH	09	UMP A24 2481	V.V. Anisovich, A.V. Sarantsev	
UEHARA	09	PR D79 052009	S. Uehara <i>et al.</i>	(BELLE Collab.)
AMELIN	06	PAN 69 690	D.V. Amelin <i>et al.</i>	(VES Collab.)
Translated from YAF 69 715.				
BINON	05	PAN 68 960	F. Binon <i>et al.</i>	
AMELIN	00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
BARBERIS	00F	PL B484 198	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ALDE	98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
Also		PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
Translated from YAF 62 446.				
MARTIN	98	PR C57 3492	B.R. Martin <i>et al.</i>	
MARTIN	97	PR C56 1114	B.R. Martin, G.C. Oades	(LOUC, AARH)
KLOET	96	PR D53 6120	W.M. Kloet, F. Myhrer	(RUTG, NORD)
OAKDEN	94	NP A574 731	M.N. Oakden, M.R. Pennington	(DURH)
ALDE	90	PL B241 600	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP+)
OEST	90	ZPHY C47 343	T. Oest <i>et al.</i>	(JADE Collab.)
ALDE	87	PL B198 286	D.M. Alde <i>et al.</i>	(LANL, BRUX, SERP, LAPP)
AUGUSTIN	87	ZPHY C36 369	J.E. Augustin <i>et al.</i>	(LALO, CLER, FRAS+)
BALTRUSAITIS	87	PR D35 2077	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
ALDE	86d	NP B269 485	D.M. Alde <i>et al.</i>	(BELG, LAPP, SERP, CERN+)
ALTHOFF	85b	ZPHY C29 189	M. Althoff <i>et al.</i>	(TASSO Collab.)
BINON	84b	LNC 39 41	F.G. Binon <i>et al.</i>	(SERP, BELG, LAPP)
BINON	83c	SJNP 38 723	F.G. Binon <i>et al.</i>	(SERP, BRUX+)
Translated from YAF 38 1199.				
DENNEY	83	PR D28 2726	D.L. Denney <i>et al.</i>	(IOWA, MICH)
CASON	82	PRL 48 1316	N.M. Cason <i>et al.</i>	(NDAM, ANL)
ETKIN	82b	PR D25 1786	A. Etkin <i>et al.</i>	(BNL, CUNY, TUFTS, VAND)
ALPER	80	PL 94B 422	B. Alper <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
ROZANSKA	80	NP B162 505	M. Rozanska <i>et al.</i>	(MPIM, CERN)
CORDEN	79	NP B157 250	M.J. Corden <i>et al.</i>	(BIRM, RHEL, TELA+)JP
EVANGELISTA	79b	NP B154 381	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
ANTIPOV	77	NP B119 45	Y.M. Antipov <i>et al.</i>	(SERP, GEVA)
APEL	75	PL 57B 398	W.D. Apel <i>et al.</i>	(KARLK, KARLE, PISA, SERP+)JP
BLUM	75	PL 57B 403	W. Blum <i>et al.</i>	(CERN, MPIM)JP

$\pi_2(2100)$	$J^G(J^{PC}) = 1^-(2^{-+})$
OMITTED FROM SUMMARY TABLE	
Needs confirmation.	

### $\pi_2(2100)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>2090\pm 29</math> OUR AVERAGE</b>			
$2090\pm 30$	1 AMELIN	95b	VES $36\pi^-A\rightarrow\pi^+\pi^-\pi^-A$
$2100\pm150$	2 DAUM	81b	CNTR $63,94\pi^-p\rightarrow 3\pi X$
1 From a fit to $J^{PC}=2^{-+} f_2(1270)\pi, (\pi\pi)_S\pi$ waves.			
2 From a two-resonance fit to four $2^{-0+}$ waves.			

### $\pi_2(2100)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>625\pm 50</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.		
$520\pm100$	3 AMELIN	95b	VES $36\pi^-A\rightarrow\pi^+\pi^-\pi^-A$
$651\pm 50$	4 DAUM	81b	CNTR $63,94\pi^-p\rightarrow 3\pi X$
3 From a fit to $J^{PC}=2^{-+} f_2(1270)\pi, (\pi\pi)_S\pi$ waves.			
4 From a two-resonance fit to four $2^{-0+}$ waves.			

### $\pi_2(2100)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $3\pi$	seen
$\Gamma_2$ $\rho\pi$	seen
$\Gamma_3$ $f_2(1270)\pi$	seen
$\Gamma_4$ $(\pi\pi)_S\pi$	seen

### $\pi_2(2100)$ BRANCHING RATIOS

$\Gamma(\rho\pi)/\Gamma(3\pi)$	$\Gamma_2/\Gamma_1$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.19\pm0.05</math></b>	5 DAUM	81b	CNTR	$63,94\pi^-p$

$\Gamma(f_2(1270)\pi)/\Gamma(3\pi)$	$\Gamma_3/\Gamma_1$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.36\pm0.09</math></b>	5 DAUM	81b	CNTR	$63,94\pi^-p$

$\Gamma((\pi\pi)_S\pi)/\Gamma(3\pi)$	$\Gamma_4/\Gamma_1$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.45\pm0.07</math></b>	5 DAUM	81b	CNTR	$63,94\pi^-p$

### D-wave/S-wave RATIO FOR $\pi_2(2100)\rightarrow f_2(1270)\pi$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.39\pm0.23</math></b>	5 DAUM	81b	CNTR $63,94\pi^-p$
5 From a two-resonance fit to four $2^{-0+}$ waves.			

### $\pi_2(2100)$ REFERENCES

AMELIN	95b	PL B356 595	D.V. Amelin <i>et al.</i>	(SERP, TBIL)
DAUM	81b	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)

$f_0(2100)$	$J^G(J^{PC}) = 0^+(0^{++})$
OMITTED FROM SUMMARY TABLE	
Needs confirmation.	

### $f_0(2100)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2101\pm 7</math> OUR AVERAGE</b>				
$2081\pm13^{+24}_{-36}$	5.5k	1 ABLIKIM	13N	BES3 $e^+e^-\rightarrow J/\psi\rightarrow\gamma\eta\eta$
$2102\pm13$		2 ANISOVICH	00J	SPEC $2.0\overline{p}p\rightarrow\eta\pi^0\pi^0, \pi^0\pi^0, \eta\eta, \eta\eta', \pi^+\pi^-$
$2090\pm30$		BAI	00A	BES $J/\psi\rightarrow\gamma(\pi^+\pi^-\pi^+\pi^-)$
$2105\pm10$		ANISOVICH	99K	SPEC $0.6\text{--}1.94\overline{p}p\rightarrow\eta\eta, \eta\eta'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2090\pm10\pm 6$	529	3,4 DOBBS	15	$J/\psi\rightarrow\gamma\pi^+\pi^-$
$2099\pm17\pm 8$	283	3,4 DOBBS	15	$\psi(2S)\rightarrow\gamma\pi^+\pi^-$
$2105\pm 8$	80k	5 UMAN	06	E835 $5.2\overline{p}p\rightarrow\eta\eta\pi^0$
$\sim 2104$		BUGG	95	$J/\psi\rightarrow\gamma\pi^+\pi^-\pi^+\pi^-$
$\sim 2122$		HASAN	94	RVUE $\overline{p}p\rightarrow\pi\pi$

- 1 From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.
- 2 Includes the data of ANISOVICH 00b indicating to exotic decay pattern.
- 3 Using CLEO-c data but not authored by the CLEO Collaboration.
- 4 From a fit to a Breit-Wigner line shape with fixed  $\Gamma=209$  MeV.
- 5 Statistical error only.

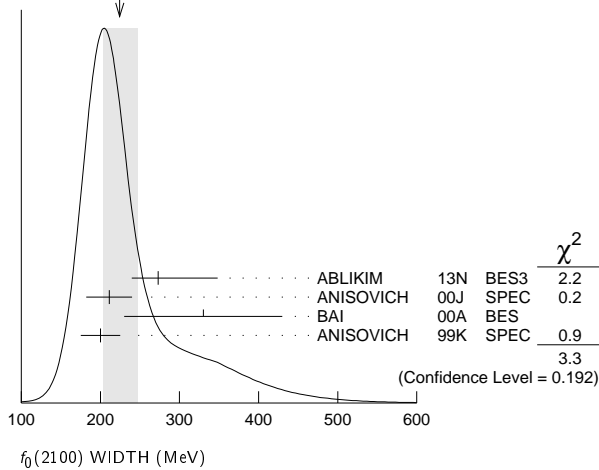
### $f_0(2100)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>224^{+23}_{-21}</math> OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
$273^{+27+70}_{-24-23}$	5.5k	6 ABLIKIM	13N	BES3 $e^+e^-\rightarrow J/\psi\rightarrow\gamma\eta\eta$
$211\pm 29$		7 ANISOVICH	00J	SPEC $2.0\overline{p}p\rightarrow\eta\pi^0\pi^0, \pi^0\pi^0, \eta\eta, \eta\eta', \pi^+\pi^-$
$330\pm100$		BAI	00A	BES $J/\psi\rightarrow\gamma(\pi^+\pi^-\pi^+\pi^-)$
$200\pm 25$		ANISOVICH	99K	SPEC $0.6\text{--}1.94\overline{p}p\rightarrow\eta\eta, \eta\eta'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$236\pm 14$	80k	8 UMAN	06	E835 $5.2\overline{p}p\rightarrow\eta\eta\pi^0$
$\sim 203$		BUGG	95	$J/\psi\rightarrow\gamma\pi^+\pi^-\pi^+\pi^-$
$\sim 273$		HASAN	94	RVUE $\overline{p}p\rightarrow\pi\pi$

- 6 From partial wave analysis including all possible combinations of  $0^{++}$ ,  $2^{++}$ , and  $4^{++}$  resonances.
- 7 Includes the data of ANISOVICH 00b indicating to exotic decay pattern.
- 8 Statistical error only.

See key on page 601

## Meson Particle Listings

 $f_0(2100)$ ,  $f_2(2150)$ WEIGHTED AVERAGE  
224±23-21 (Error scaled by 1.3) $f_0(2100)$  REFERENCES

DOBBS 15	PR D91 052006	S. Dobbs <i>et al.</i>	(NWES)
ABLIKIM 13N	PR D87 092009	Ablikim M. <i>et al.</i>	(BES III Collab.)
UMAN 06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
ANISOVICH 00B	NP A662 319	A.V. Anisovich <i>et al.</i>	
ANISOVICH 00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
BAI 00A	PL B472 207	J.Z. Bai <i>et al.</i>	(BES Collab.)
ANISOVICH 99K	PL B468 309	A.V. Anisovich <i>et al.</i>	
BUGG 95	PL B353 378	D.V. Bugg <i>et al.</i>	(LOQM, PNPI, WASH)
HASAN 94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)

 $f_2(2150)$ 

$$I^G(J^{PC}) = 0^+(2^{++})$$

OMITTED FROM SUMMARY TABLE  
This entry was previously called  $T_0$ . $f_2(2150)$  MASS $f_2(2150)$  MASS, COMBINED MODES (MeV)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2157±12 OUR AVERAGE</b>	Includes data from the 2 datablocks that follow this one.			
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2170±6	80k	<sup>1</sup> UMAN	06 E835	5.2 $\bar{p}p \rightarrow \eta\eta\pi^0$
<sup>1</sup> Statistical error only.				

 $\eta\eta$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.			

**2157±12 OUR AVERAGE**

2151±16	BARBERIS 00E	450 $pp \rightarrow p_f \eta \eta p_s$	
2175±20	PROKOSHKIN 95D	GAM4 300 $\pi^- N \rightarrow \pi^- N 2\eta$ , 450 $pp \rightarrow pp 2\eta$	
2130±35	SINGOVSKI 94	GAM4 450 $pp \rightarrow pp 2\eta$	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2140±30	<sup>2</sup> ABELE 99B	CBAR	
2104±20	<sup>3</sup> ARMSTRONG 93c	E760 $\bar{p}p \rightarrow \pi^0 \eta \eta \rightarrow 6\gamma$	
<sup>2</sup> Spin not determined.			
<sup>3</sup> No $J^{PC}$ determination.			

 $\eta\pi\pi$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.				

• • • We do not use the following data for averages, fits, limits, etc. • • •				
2135±20±45	<sup>4</sup> ADOMEIT 96	CBAR 0	1.94 $\bar{p}p \rightarrow \eta 3\pi^0$	
<sup>4</sup> ANISOVICH 00E recommends to withdraw ADOMEIT 96 that assumed a single $J^P = 2^+$ resonance.				

 $\bar{p}p \rightarrow \pi\pi$ 

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
~2090	<sup>5</sup> OAKDEN 94	RVUE	0.36–1.55 $\bar{p}p \rightarrow \pi\pi$
~2120	<sup>6</sup> OAKDEN 94	RVUE	0.36–1.55 $\bar{p}p \rightarrow \pi\pi$
~2170	<sup>7</sup> MARTIN 80b	RVUE	
~2150	<sup>7</sup> MARTIN 80c	RVUE	
~2150	<sup>8</sup> DULUDE 78b	OSPK	1–2 $\bar{p}p \rightarrow \pi^0 \pi^0$

<sup>5</sup> OAKDEN 94 makes an amplitude analysis of LEAR data on  $\bar{p}p \rightarrow \pi\pi$  using a method based on Barrelet zeros. This is solution A. The amplitude analysis of HASAN 94 includes earlier data as well, and assume that the data can be parametrized in terms of towers of nearly degenerate resonances on the leading Regge trajectory. See also KLOET 96 and MARTIN 97 who make related analyses.

<sup>6</sup> From solution B of amplitude analysis of data on  $\bar{p}p \rightarrow \pi\pi$ .

<sup>7</sup>  $I(J^P) = 0(2^+)$  from simultaneous analysis of  $p\bar{p} \rightarrow \pi^- \pi^+ \pi^0$  and  $\pi^0 \pi^0$ .

<sup>8</sup>  $I^G(J^P) = 0^+(2^+)$  from partial-wave amplitude analysis.

S-CHANNEL  $\bar{p}p$ ,  $\bar{N}N$  or  $\bar{K}K$ 

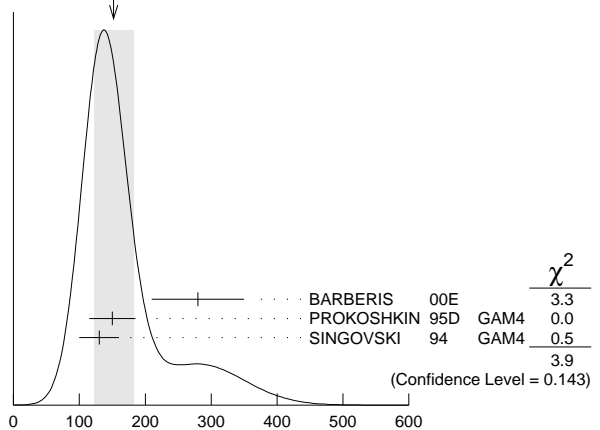
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2139± <sup>8</sup> <sub>9</sub>	<sup>9</sup> EVANGELIS...	97 SPEC		0.6–2.4 $\bar{p}p \rightarrow K_S^0 K_S^0$
~2190	<sup>9</sup> CUTTS	78B CNTR		0.97–3 $\bar{p}p \rightarrow \bar{N}N$
2155±15	<sup>9,10</sup> COUPLAND	77 CNTR 0		0.7–2.4 $\bar{p}p \rightarrow \bar{p}p$
2193±2	<sup>9,11</sup> ALSPECTOR	73 CNTR		$\bar{p}p$ S channel
<sup>9</sup> Isospins 0 and 1 not separated.				
<sup>10</sup> From a fit to the total elastic cross section.				
<sup>11</sup> Referred to as $T$ or $T$ region by ALSPECTOR 73.				

 $K\bar{K}$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2200±13	VLADIMIRSK..06	SPEC	40 $\pi^- p \rightarrow K_S^0 K_S^0 n$
2150±20	ABLIKIM 04E	BES2	$J/\psi \rightarrow \omega K^+ K^-$
2130±35	BARBERIS 99	OMEG	450 $pp \rightarrow p_s p_f K^+ K^-$

 $f_2(2150)$  WIDTH $f_2(2150)$  WIDTH, COMBINED MODES (MeV)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>152±30 OUR AVERAGE</b>	Includes data from the 2 datablocks that follow this one. Error includes scale factor of 1.4. See the ideogram below.			
• • • We do not use the following data for averages, fits, limits, etc. • • •				
182±11	80k	<sup>12</sup> UMAN	06 E835	5.2 $\bar{p}p \rightarrow \eta\eta\pi^0$
<sup>12</sup> Statistical error only.				

WEIGHTED AVERAGE  
152±30 (Error scaled by 1.4) $f_2(2150)$  WIDTH, COMBINED MODES (MeV) $\eta\eta$  MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.			

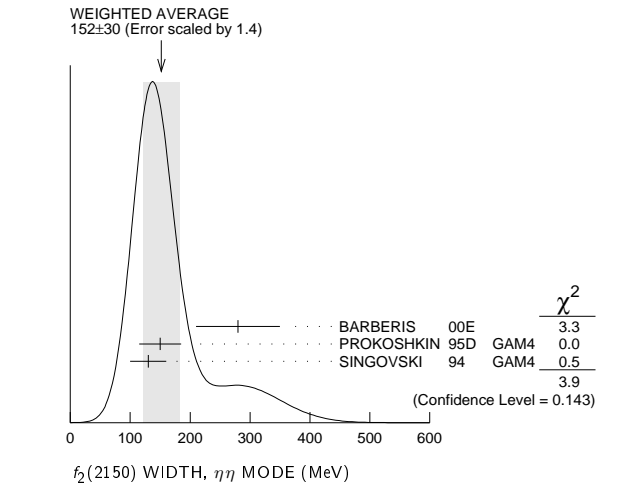
**152±30 OUR AVERAGE** Error includes scale factor of 1.4. See the ideogram below.

280±70	BARBERIS 00E	450 $pp \rightarrow p_f \eta \eta p_s$	
150±35	PROKOSHKIN 95D	GAM4 300 $\pi^- N \rightarrow \pi^- N 2\eta$ , 450 $pp \rightarrow pp 2\eta$	
130±30	SINGOVSKI 94	GAM4 450 $pp \rightarrow pp 2\eta$	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
310±50	<sup>13</sup> ABELE 99B	CBAR	
203±10	<sup>14</sup> ARMSTRONG 93c	E760	$\bar{p}p \rightarrow \pi^0 \eta \eta \rightarrow 6\gamma$

# Meson Particle Listings

## $f_2(2150)$ , $\rho(2150)$

<sup>13</sup> Spin not determined.  
<sup>14</sup> No  $J^{PC}$  determination.



### $\eta\pi\pi$ MODE

VALUE (MeV) DOCUMENT ID TECN CHG COMMENT  
The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •  
250±25±45 <sup>15</sup> ADOMEIT 96 CBAR 0 1.94  $\overline{p}p \rightarrow \eta 3\pi^0$   
<sup>15</sup> ANISOVICH 00E recommends to withdraw ADOMEIT 96 that assumed a single  $J^P = 2^+$  resonance.

### $\overline{p}p \rightarrow \pi\pi$

VALUE (MeV) DOCUMENT ID TECN COMMENT  
**250 OUR ESTIMATE**  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
~ 70 <sup>16</sup> OAKDEN 94 RVUE 0.36–1.55  $\overline{p}p \rightarrow \pi\pi$   
~ 250 <sup>17</sup> MARTIN 80B RVUE  
~ 250 <sup>17</sup> MARTIN 80C RVUE  
~ 250 <sup>18</sup> DULUDE 78B OSPK 1–2  $\overline{p}p \rightarrow \pi^0\pi^0$

<sup>16</sup> See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J = 3$  to be important but not significantly resonant.  
<sup>17</sup>  $I(J^P) = 0(2^+)$  from simultaneous analysis of  $p\overline{p} \rightarrow \pi^-\pi^+$  and  $\pi^0\pi^0$ .  
<sup>18</sup>  $I^G(J^P) = 0^+(2^+)$  from partial-wave amplitude analysis.

### S-CHANNEL $\overline{p}p$ , $\overline{N}N$ or $\overline{K}K$

VALUE (MeV) DOCUMENT ID TECN CHG COMMENT  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
56+31-16 <sup>19</sup> EVANGELIS... 97 SPEC 0.6–2.4  $\overline{p}p \rightarrow K_S^0 K_S^0$   
135±75 <sup>20,21</sup> COUPLAND 77 CNTR 0 0.7–2.4  $\overline{p}p \rightarrow \overline{p}p$   
98±8 <sup>21</sup> ALSPECTOR 73 CNTR  $\overline{p}p$  S channel  
<sup>19</sup> Isospin 0 and 2 not separated.  
<sup>20</sup> From a fit to the total elastic cross section.  
<sup>21</sup> Isospins 0 and 1 not separated.

### $K\overline{K}$ MODE

VALUE (MeV) DOCUMENT ID TECN COMMENT  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
91±62 VLADIMIRSK...06 SPEC 40  $\pi^-p \rightarrow K_S^0 K_S^0 n$   
150±30 ABLIKIM 04E BES2  $J/\psi \rightarrow \omega K^+ K^-$   
270±50 BARBERIS 99 OMEG 450  $pp \rightarrow p_S p_f K^+ K^-$

### $f_2(2150)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi\pi$	
$\Gamma_2$ $\eta\eta$	seen
$\Gamma_3$ $K\overline{K}$	seen
$\Gamma_4$ $f_2(1270)\eta$	seen
$\Gamma_5$ $a_2(1320)\pi$	seen
$\Gamma_6$ $\rho\overline{\rho}$	seen

### $f_2(2150)$ BRANCHING RATIOS

#### $\Gamma(K\overline{K})/\Gamma(\eta\eta)$

VALUE CL% DOCUMENT ID TECN COMMENT  
**1.28±0.23** BARBERIS 00E 450  $pp \rightarrow p_f \eta \eta p_S$   
• • • We do not use the following data for averages, fits, limits, etc. • • •  
<0.1 95 <sup>22</sup> PROKOSHKIN 95D GAM4 300  $\pi^-N \rightarrow \pi^- N 2\eta$ ,  
450  $pp \rightarrow pp 2\eta$   
<sup>22</sup> Using data from ARMSTRONG 89D.

#### $\Gamma(\pi\pi)/\Gamma(\eta\eta)$

VALUE CL% DOCUMENT ID TECN COMMENT  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
<0.33 95 <sup>23</sup> PROKOSHKIN 95D GAM4 300  $\pi^-N \rightarrow \pi^- N 2\eta$ ,  
450  $pp \rightarrow pp 2\eta$   
<sup>23</sup> Derived from a  $\pi^0\pi^0/\eta\eta$  limit.

#### $\Gamma(f_2(1270)\eta)/\Gamma(a_2(1320)\pi)$

VALUE DOCUMENT ID TECN COMMENT  
**0.79±0.11** <sup>24</sup> ADOMEIT 96 CBAR 1.94  $\overline{p}p \rightarrow \eta 3\pi^0$   
<sup>24</sup> Using  $B(a_2(1320) \rightarrow \eta\pi) = 0.145$

#### $\Gamma(\rho\overline{\rho})/\Gamma_{\text{total}}$

VALUE EVTS DOCUMENT ID TECN COMMENT  
**seen** 73 ALEXANDER 10 CLEO  $\psi(2S) \rightarrow \gamma\rho\overline{\rho}$

### $f_2(2150)$ REFERENCES

ALEXANDER 89D	PR D82 092002	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
UMAN 06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
VLADIMIRSK...06	PAN 69 493	V.V. Vladimirovsky <i>et al.</i>	(ITEP, Moscow)
	Translated from YAF 69 515.		
ABLIKIM 04E	PL B603 138	M. Ablikim <i>et al.</i>	(BES Collab.)
ANISOVICH 00E	PL B477 19	A.V. Anisovich <i>et al.</i>	(DURH)
BARBERIS 00E	PL B479 59	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ABELE 99B	EPJ C8 67	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BARBERIS 99	PL B453 305	D. Barberis <i>et al.</i>	(Omega Expt.)
EVANGELIS... 97	PR D56 3803	C. Evangelista <i>et al.</i>	(LEAR Collab.)
MARTIN 97	PR C56 1114	B.R. Martin, G.C. Oades	(LOUC, AARH)
ADOMEIT 96	ZPHY C71 227	J. Adomeit <i>et al.</i>	(Crystal Barrel Collab.)
KLOET 96	PR D53 6120	W.M. Kloet, F. Myhrer	(RUTG, NORD)
PROKOSHKIN 95D	SPD 40 495	Y.D. Prokoshkin	(SERP)IGJPC
	Translated from DANS 344 469.		
HASAN 94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)
OAKDEN 94	NP A574 731	M.N. Oakden, M.R. Pennington	(DURH)
SINGOVSKI 94	NC 107A 1911	A.V. Singovsky	(SERP)
ARMSTRONG 93C	PL B307 394	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
ARMSTRONG 89D	PL B227 186	T.A. Armstrong, M. Benayoun	(ATHU, BARI, BIRM+)
MARTIN 80B	NP B176 355	B.R. Martin, D. Morgan	(LOUC, RHEL)JP
MARTIN 80C	NP B169 216	A.D. Martin, M.R. Pennington	(DURH)JP
CUTTS 78B	PR D17 16	D. Cutts <i>et al.</i>	(STON, WISC)
DULUDE 78B	PL 79B 335	R.S. Dulude <i>et al.</i>	(BROW, MIT, BARI)JP
COUPLAND 77	PL 71B 460	M. Coupland <i>et al.</i>	(LOQM, RHEL)
ALSPECTOR 73	PRL 30 511	J. Alspector <i>et al.</i>	(RUTG, UPNJ)

## $\rho(2150)$

$$I^G(J^{PC}) = 1^+(1^-)$$

### OMITTED FROM SUMMARY TABLE

This entry was previously called  $T_1(2190)$ . See our mini-review under the  $\rho(1700)$ .

### $\rho(2150)$ MASS

### $e^+e^-$ PRODUCED

VALUE (MeV) DOCUMENT ID TECN COMMENT  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
2254±22 <sup>1</sup> LEES 12G BABR  $e^+e^- \rightarrow \pi^+\pi^-\gamma$   
2150±40±50 AUBERT 07AU BABR 10.6  $e^+e^- \rightarrow f_1(1285)\pi^+\pi^-\gamma$   
1990±80 AUBERT 07AU BABR 10.6  $e^+e^- \rightarrow \eta'\pi^+\pi^-\gamma$   
2153±37 BIAGINI 91 RVUE  $e^+e^- \rightarrow \pi^+\pi^-, K^+K^-$   
2110±50 <sup>2</sup> CLEGG 90 RVUE  $e^+e^- \rightarrow 3(\pi^+\pi^-), 2(\pi^+\pi^-\pi^0)$

### $\overline{p}p \rightarrow \pi\pi$

VALUE (MeV) DOCUMENT ID TECN COMMENT  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
~ 2191 HASAN 94 RVUE  $\overline{p}p \rightarrow \pi\pi$   
~ 2070 <sup>3</sup> OAKDEN 94 RVUE 0.36–1.55  $\overline{p}p \rightarrow \pi\pi$   
~ 2170 <sup>4</sup> MARTIN 80B RVUE  
~ 2100 <sup>4</sup> MARTIN 80C RVUE

### S-CHANNEL $\overline{N}N$

VALUE (MeV) DOCUMENT ID TECN COMMENT  
• • • We do not use the following data for averages, fits, limits, etc. • • •  
2110±35 <sup>5</sup> ANISOVICH 02 SPEC 0.6–1.9  $\overline{p}p \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$   
~ 2190 <sup>6</sup> CUTTS 78B CNTR 0.97–3  $\overline{p}p \rightarrow \overline{N}N$   
2155±15 <sup>6,7</sup> COUPLAND 77 CNTR 0.7–2.4  $\overline{p}p \rightarrow \overline{p}p$   
2193±2 <sup>6,8</sup> ALSPECTOR 73 CNTR  $\overline{p}p$  S channel  
2190±10 <sup>9</sup> ABRAMS 70 CNTR S channel  $\overline{p}N$

See key on page 601

## Meson Particle Listings

 $\rho(2150), \phi(2170)$  $\pi^- p \rightarrow \omega \pi^0 n$ 

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2155±21 OUR AVERAGE</b>			
2140±30	ALDE	95	GAM2 38 $\pi^- p \rightarrow \omega \pi^0 n$
2170±30	ALDE	92c	GAM4 100 $\pi^- p \rightarrow \omega \pi^0 n$

<sup>1</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>2</sup> Includes ATKINSON 85.

<sup>3</sup> See however KLOET 96 who fit  $\pi^+ \pi^-$  only and find waves only up to  $J = 3$  to be important but not significantly resonant.

<sup>4</sup>  $I(J^P) = 1(1^-)$  from simultaneous analysis of  $p\bar{p} \rightarrow \pi^- \pi^+$  and  $\pi^0 \pi^0$ .

<sup>5</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

<sup>6</sup> Isospins 0 and 1 not separated.

<sup>7</sup> From a fit to the total elastic cross section.

<sup>8</sup> Referred to as  $T$  or  $T$  region by ALSPECTOR 73.

<sup>9</sup> Seen as bump in  $l = 1$  state. See also COOPER 68. PEASLEE 75 confirm  $\bar{p}p$  results of ABRAMS 70, no narrow structure.

 $\rho(2150)$  WIDTH $e^+ e^-$  PRODUCED

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
109± 76	<sup>10</sup> LEES	12G	BABR $e^+ e^- \rightarrow \pi^+ \pi^- \gamma$
350± 40±50	AUBERT	07AU	BABR 10.6 $e^+ e^- \rightarrow f_1(1285) \pi^+ \pi^- \gamma$
310±140	AUBERT	07AU	BABR 10.6 $e^+ e^- \rightarrow \eta' \pi^+ \pi^- \gamma$
389± 79	BIAGINI	91	RVUE $e^+ e^- \rightarrow \pi^+ \pi^-, K^+ K^-$
410±100	<sup>11</sup> CLEGG	90	RVUE $e^+ e^- \rightarrow 3(\pi^+ \pi^-), 2(\pi^+ \pi^- \pi^0)$

 $\bar{p}p \rightarrow \pi\pi$ 

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
~ 296	HASAN	94	RVUE $\bar{p}p \rightarrow \pi\pi$
~ 40	<sup>12</sup> OAKDEN	94	RVUE 0.36–1.55 $\bar{p}p \rightarrow \pi\pi$
~ 250	<sup>13</sup> MARTIN	80B	RVUE
~ 200	<sup>13</sup> MARTIN	80C	RVUE

S-CHANNEL  $\bar{N}N$ 

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
230±50	<sup>14</sup> ANISOVICH	02	SPEC 0.6–1.9 $p\bar{p} \rightarrow \omega \pi^0, \omega \eta \pi^0, \pi^+ \pi^-$
135±75	<sup>15,16</sup> COUPLAND	77	CNTR 0.7–2.4 $\bar{p}p \rightarrow \bar{p}p$
98± 8	<sup>16</sup> ALSPECTOR	73	CNTR $\bar{p}p$ S channel
~ 85	<sup>17</sup> ABRAMS	70	CNTR S channel $\bar{p}N$

 $\pi^- p \rightarrow \omega \pi^0 n$ 

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>320±70</b>	ALDE	95	GAM2 38 $\pi^- p \rightarrow \omega \pi^0 n$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
~ 300	ALDE	92c	GAM4 100 $\pi^- p \rightarrow \omega \pi^0 n$

<sup>10</sup> Using the GOUNARIS 68 parametrization of the pion form factor leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>11</sup> Includes ATKINSON 85.

<sup>12</sup> See however KLOET 96 who fit  $\pi^+ \pi^-$  only and find waves only up to  $J = 3$  to be important but not significantly resonant.

<sup>13</sup>  $I(J^P) = 1(1^-)$  from simultaneous analysis of  $p\bar{p} \rightarrow \pi^- \pi^+$  and  $\pi^0 \pi^0$ .

<sup>14</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

<sup>15</sup> From a fit to the total elastic cross section.

<sup>16</sup> Isospins 0 and 1 not separated.

<sup>17</sup> Seen as bump in  $l = 1$  state. See also COOPER 68. PEASLEE 75 confirm  $\bar{p}p$  results of ABRAMS 70, no narrow structure.

 $\rho(2150)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $e^+ e^-$	
$\Gamma_2$ $\pi^+ \pi^-$	seen
$\Gamma_3$ $K^+ K^-$	seen
$\Gamma_4$ $3(\pi^+ \pi^-)$	seen
$\Gamma_5$ $2(\pi^+ \pi^- \pi^0)$	seen
$\Gamma_6$ $\eta' \pi^+ \pi^-$	seen
$\Gamma_7$ $f_1(1285) \pi^+ \pi^-$	seen
$\Gamma_8$ $\omega \pi^0$	seen
$\Gamma_9$ $\omega \pi^0 \eta$	seen
$\Gamma_{10}$ $\rho\bar{\rho}$	

 $\rho(2150) \Gamma(i) \Gamma(e^+ e^-) / \Gamma^2(\text{total})$ 

$$\Gamma(f_1(1285) \pi^+ \pi^-) / \Gamma_{\text{total}} \times \Gamma(e^+ e^-) / \Gamma_{\text{total}} \quad \Gamma_7 / \Gamma \times \Gamma_1 / \Gamma$$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.1±0.6±0.5</b>	<sup>18</sup> AUBERT	07AU	BABR 10.6 $e^+ e^- \rightarrow f_1(1285) \pi^+ \pi^- \gamma$

<sup>18</sup> Calculated by us from the reported value of cross section at the peak.

$$\Gamma(\eta' \pi^+ \pi^-) / \Gamma_{\text{total}} \times \Gamma(e^+ e^-) / \Gamma_{\text{total}} \quad \Gamma_6 / \Gamma \times \Gamma_1 / \Gamma$$

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			

4.9±1.9	<sup>19</sup> AUBERT	07AU	BABR 10.6 $e^+ e^- \rightarrow \eta' \pi^+ \pi^- \gamma$
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<sup>19</sup> Calculated by us from the reported value of cross section at the peak.

 $\rho(2150)$  REFERENCES

LEES	12G	PR D86 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
AUBERT	07AU	PR D76 092005	B. Aubert <i>et al.</i>	(BABAR Collab.)
ANISOVICH	02	PL B542 8	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01D	PL B508 6	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01E	PL B513 281	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
KLOET	96	PR D53 6120	W.M. Kloeit, F. Myhrer	(RUTG, NORD)
ALDE	95	ZPHY C66 379	D.M. Alde <i>et al.</i>	(GAMS Collab.) JP
HASAN	94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)
OAKDEN	94	NP A574 731	M.N. Oakden, M.R. Pennington	(DURH)
ALDE	92C	ZPHY C54 553	D.M. Alde <i>et al.</i>	(BELG, SERP, KEK, LANL+)
BIAGINI	91	NC 104A 363	M.E. Biagini <i>et al.</i>	(FRAS, PRAG)
CLEGG	90	ZPHY C45 677	A.B. Clegg, A. Donnachie	(LANC, MCHS)
ATKINSON	85	ZPHY C29 333	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
MARTIN	80B	NP B176 355	B.R. Martin, D. Morgan	(LOUC, RHEL) JP
MARTIN	80C	NP B169 216	A.D. Martin, M.R. Pennington	(DURH) JP
CUTTS	78B	PR D17 16	D. Cutts <i>et al.</i>	(STON, WISC)
COUPLAND	77	PL 71B 460	M. Coupland <i>et al.</i>	(LOQM, RHEL)
PEASLEE	75	PL 57B 189	D.C. Peaslee <i>et al.</i>	(CANB, BARI, BROW+)
ALSPECTOR	73	PRL 30 511	J. Alspector <i>et al.</i>	(RUTG, UPNJ)
ABRAMS	70	PR D1 1917	R.J. Abrams <i>et al.</i>	(BNL)
COOPER	68	PRL 20 1059	W.A. Cooper <i>et al.</i>	(ANL)
GOUNARIS	68	PRL 21 244	G.J. Gounaris, J.J. Sakurai	

 $\phi(2170)$ 

$$I_G(J^P C) = 0^-(1^- -)$$

Observed by AUBERT, BE 06D in the initial-state radiation process  
 $e^+ e^- \rightarrow \phi f_0(980) \gamma$ .

 $\phi(2170)$  MASS

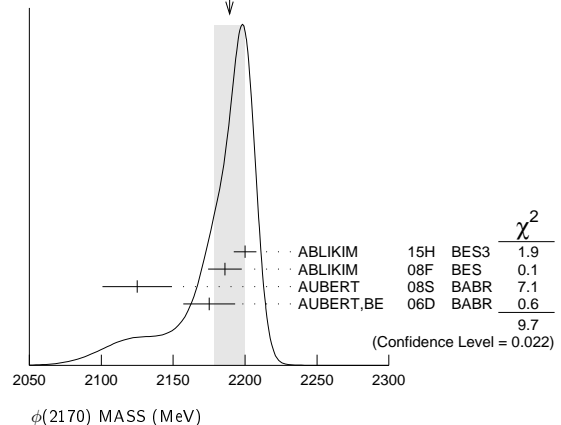
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2189±11 OUR AVERAGE</b>				Error includes scale factor of 1.8. See the ideogram below.
2200± 6± 5	471	ABLIKIM	15H	BES3 $J/\psi \rightarrow \eta \phi \pi^+ \pi^-$
2186±10± 6	52	ABLIKIM	08F	BES $J/\psi \rightarrow \eta \phi f_0(980)$
2125±22±10	483	AUBERT	08S	BABR 10.6 $e^+ e^- \rightarrow \phi \eta \gamma$
2175±10±15	201	<sup>1</sup> AUBERT, BE	06D	BABR 10.6 $e^+ e^- \rightarrow K^+ K^- \pi \pi \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2079±13± <sup>79</sup> <sub>28</sub>	4.8k	<sup>2</sup> SHEN	09	BELL 10.6 $e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$
2192±14	116	<sup>3</sup> AUBERT	07AK	BABR 10.6 $e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$
2169±20	149	<sup>3</sup> AUBERT	07AK	BABR 10.6 $e^+ e^- \rightarrow K^+ K^- \pi^0 \pi^0 \gamma$

<sup>1</sup> From the  $\phi f_0(980)$  component.

<sup>2</sup> From a fit with two incoherent Breit-Wigners.

<sup>3</sup> From the  $K^+ K^- \phi(980)$  component.

WEIGHTED AVERAGE  
 2189±11 (Error scaled by 1.8)



# Meson Particle Listings

## $\phi(2170)$ , $f_0(2200)$ , $f_J(2220)$

### $\phi(2170)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>79±14 OUR AVERAGE</b>				
104±15±15	471	ABLIKIM	15H BES3	$J/\psi \rightarrow \eta \phi \pi^+ \pi^-$
65±23±17	52	ABLIKIM	08F BES	$J/\psi \rightarrow \eta \phi f_0(980)$
61±50±13	483	AUBERT	08S BABR	10.6 $e^+ e^- \rightarrow \phi \eta \gamma$
58±16±20	201	<sup>4</sup> AUBERT,BE	06D BABR	10.6 $e^+ e^- \rightarrow K^+ K^- \pi \pi \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
192±23+ <sup>25</sup> <sub>-61</sub>	4.8k	<sup>5</sup> SHEN	09 BELL	10.6 $e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$
71±21	116	<sup>6</sup> AUBERT	07AK BABR	10.6 $e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$
102±27	149	AUBERT	07AK BABR	10.6 $e^+ e^- \rightarrow K^+ K^- \pi^0 \pi^0 \gamma$
<sup>4</sup> From the $\phi f_0(980)$ component.				
<sup>5</sup> From a fit with two incoherent Breit-Wigners.				
<sup>6</sup> From the $K^+ K^- f_0(980)$ component.				

### $\phi(2170)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $e^+ e^-$	seen
$\Gamma_2$ $\phi \eta$	
$\Gamma_3$ $\phi \pi \pi$	
$\Gamma_4$ $\phi f_0(980)$	seen
$\Gamma_5$ $K^+ K^- \pi^+ \pi^-$	
$\Gamma_6$ $K^+ K^- f_0(980) \rightarrow K^+ K^- \pi^+ \pi^-$	seen
$\Gamma_7$ $K^+ K^- \pi^0 \pi^0$	
$\Gamma_8$ $K^+ K^- f_0(980) \rightarrow K^+ K^- \pi^0 \pi^0$	seen
$\Gamma_9$ $K^{*0} K^\pm \pi^\mp$	not seen
$\Gamma_{10}$ $K^*(892)^0 \bar{K}^*(892)^0$	not seen

### $\phi(2170)$ $\Gamma(i)\Gamma(e^+ e^-)/\Gamma(\text{total})$

$\Gamma(\phi\eta) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_2\Gamma_1/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.7±0.7±1.3	483	AUBERT	08S BABR	10.6 $e^+e^- \rightarrow \phi\eta\gamma$

$\Gamma(\phi f_0(980)) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_4 \Gamma_1/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.5±0.8±0.4</b>	201	<sup>7</sup> AUBERT,BE	06D BABR	10.6 $e^+ e^- \rightarrow K^+ K^- \pi \pi \gamma$	
<sup>7</sup> From the $\phi f_0(980)$ component.					

### $\phi(2170)$ $\Gamma(i)\Gamma(e^+ e^-)/\Gamma^2(\text{total})$

$\Gamma(\phi\pi\pi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma \times \Gamma_1/\Gamma$
VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.65 \pm 0.15 \pm 0.18$	4.8k	<sup>8</sup> SHEN	09 BELL	$10.6 e^+e^- \rightarrow K^+K^-\pi^+\pi^-\gamma$	
<sup>8</sup> Multiplied by 3/2 to take into account the $\phi\pi^0\pi^0$ mode. Using $B(\phi \rightarrow K^+K^-) = (49.2 \pm 0.6)\%$ .					

### $\phi(2170)$ BRANCHING RATIOS

$\Gamma(K^+ K^- f_0(980) \rightarrow K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	07AK BABR	$10.6\text{ e}^+ \text{e}^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$	
$\Gamma(K^+ K^- f_0(980) \rightarrow K^+ K^- \pi^0 \pi^0)/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	07AK BABR	$10.6\text{ e}^+ \text{e}^- \rightarrow K^+ K^- \pi^0 \pi^0 \gamma$	
$\Gamma(K^{*0} K^\pm \pi^\mp)/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
not seen	AUBERT	07AK BABR	$10.6\text{ GeV e}^+ \text{e}^-$	
$\Gamma(K^*(892)^0 \bar{K}^*(892)^0)/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
not seen	ABLIKIM	10C BES2	$J/\psi \rightarrow \eta K^+ \pi^- K^- \pi^+$	

### $\phi(2170)$ REFERENCES

ABLIKIM	15H	PR D91 052017	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	10C	PL B685 27	M. Ablikim <i>et al.</i>	(BES II Collab.)
SHEN	09	PR D80 031101	C.P. Shen <i>et al.</i>	(BELLE Collab.)
ABLIKIM	08F	PRL 100 102003	M. Ablikim <i>et al.</i>	(BES Collab.)
AUBERT	08S	PR D77 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AK	PR D76 012008	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,BE	06D	PR D74 091103	B. Aubert <i>et al.</i>	(BABAR Collab.)

## $f_0(2200)$

$$J^G(J^{PC}) = 0^+(0^{++})$$

OMITTED FROM SUMMARY TABLE

Seen in  $K_S^0 K_S^0$  (AUGUSTIN 88),  $K^+ K^-$  (ABLIKIM 05Q) and  $\eta \eta$  (BINON 05) system. Not seen in  $\mathcal{T}(1S)$  radiative decays (BARU 89).

### $f_0(2200)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2189±13 OUR AVERAGE</b>				
2170±20+ <sup>10</sup> <sub>-15</sub>		ABLIKIM	05Q BES2	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^-$
2210±50		<sup>1</sup> BINON	05 GAMS	$33 \pi^- p \rightarrow \eta \eta n$
2197±17		<sup>2</sup> AUGUSTIN	88 DM2	$J/\psi \rightarrow \gamma K_S^0 K_S^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2206±12± 8	381	<sup>3,4</sup> DOBBS	15	$J/\psi \rightarrow \gamma K^+ K^-$
2188±17±16	203	<sup>3,4</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K^+ K^-$
~ 2122		HASAN	94 RVUE	$\bar{p} p \rightarrow \pi \pi$
~ 2321		HASAN	94 RVUE	$\bar{p} p \rightarrow \pi \pi$

<sup>1</sup> First solution, PWA is ambiguous.

<sup>2</sup> Cannot determine spin to be 0.

<sup>3</sup> Using CLEO-c data but not authored by the CLEO Collaboration.

<sup>4</sup> From a fit to a Breit-Wigner line shape with fixed  $\Gamma = 238$  MeV.

### $f_0(2200)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>238±50 OUR AVERAGE</b> Error includes scale factor of 1.2.			
220±60+ <sup>40</sup> <sub>-45</sub>	ABLIKIM	05Q BES2	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^-$
380±90	<sup>5</sup> BINON	05 GAMS	$33 \pi^- p \rightarrow \eta \eta n$
201±51	<sup>6</sup> AUGUSTIN	88 DM2	$J/\psi \rightarrow \gamma K_S^0 K_S^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
~ 273	HASAN	94 RVUE	$\bar{p} p \rightarrow \pi \pi$
~ 223	HASAN	94 RVUE	$\bar{p} p \rightarrow \pi \pi$

<sup>5</sup> First solution, PWA is ambiguous.

<sup>6</sup> Cannot determine spin to be 0.

### $f_0(2200)$ REFERENCES

DOBBS	15	PR D91 052006	S. Dobbs <i>et al.</i>	(NWES)
ABLIKIM	05Q	PR D72 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
BINON	05	PAN 68 960	F. Binon <i>et al.</i>	
Translated from YAF 68 998.				
HASAN	94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)
BARU	89	ZPHY C42 505	S.E. Baru <i>et al.</i>	(NOVO)
AUGUSTIN	88	PRL 60 2238	J.E. Augustin <i>et al.</i>	(DM2 Collab.)

## $f_J(2220)$

$$J^G(J^{PC}) = 0^+(2^{++} \text{ or } 4^{++})$$

OMITTED FROM SUMMARY TABLE

Needs confirmation. See our mini-review in the 2004 edition of this Review, PDG 04.

### $f_J(2220)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2231.1± 3.5 OUR AVERAGE</b>				
2235 ± 4 ± 6	74	BAI	96B BES	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma \pi^+ \pi^-$
2230 ± 6 ± 7	±16 46	BAI	96B BES	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma K^+ K^-$
2232 ± 8 ± 15	±15 23	BAI	96B BES	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma K_S^0 K_S^0$
2235 ± 4 ± 5	±15 32	BAI	96B BES	$e^+ e^- \rightarrow J/\psi \rightarrow \gamma p \bar{p}$
2209 ± 17 ± 15	±10	ASTON	88F LASS	$11 K^- p \rightarrow K^+ K^- \Lambda$
2230 ± 20		BOLONKIN	88 SPEC	$40 \pi^- p \rightarrow K_S^0 K_S^0 n$
2220 ± 10	41	<sup>1</sup> ALDE	86B GA24	$38\text{--}100 \pi p \rightarrow n \eta \eta'$
2230 ± 6 ± 14	±14 93	BALTRUSAIT..86D	MRK3	$e^+ e^- \rightarrow \gamma K^+ K^-$
2232 ± 7 ± 7	±7 23	BALTRUSAIT..86D	MRK3	$e^+ e^- \rightarrow \gamma K_S^0 K_S^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

2223.9± 2.5		<sup>2</sup> VLADIMIRSK..08	SPEC	$40 \pi^- p \rightarrow K_S^0 K_S^0 n + m \pi^0$
2246 ± 36		BAI	98H BES	$J/\psi \rightarrow \gamma \pi^0 \pi^0$

<sup>1</sup> ALDE 86B uses data from both the GAMS-2000 and GAMS-4000 detectors.

<sup>2</sup>  $J^{PC} = 2^{++}$ . Systematic uncertainties not evaluated



See key on page 601

## Meson Particle Listings

 $f_J(2220), \eta(2225)$  $f_J(2220)$  WIDTH

VALUE (MeV)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>23^{+8}_{-7}</math></b>			<b>OUR AVERAGE</b>		
$19^{+13}_{-11} \pm 12$	74	BAI	96B	BES	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^+\pi^-$
$20^{+20}_{-15} \pm 17$	46	BAI	96B	BES	$e^+e^- \rightarrow J/\psi \rightarrow \gamma K^+K^-$
$20^{+25}_{-16} \pm 14$	23	BAI	96B	BES	$e^+e^- \rightarrow J/\psi \rightarrow \gamma K_S^0 K_S^0$
$15^{+12}_{-9} \pm 9$	32	BAI	96B	BES	$e^+e^- \rightarrow J/\psi \rightarrow \gamma p\bar{p}$
$60^{+107}_{-57}$		ASTON	88F	LASS	$11 K^-p \rightarrow K^+K^-A$
$80 \pm 30$		BOLONKIN	88	SPEC	$40 \pi^-p \rightarrow K_S^0 K_S^0 n$
$26^{+20}_{-16} \pm 17$	93	BALTRUSAIT...86D	MRK3		$e^+e^- \rightarrow \gamma K^+K^-$
$18^{+23}_{-15} \pm 10$	23	BALTRUSAIT...86D	MRK3		$e^+e^- \rightarrow \gamma K_S^0 K_S^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$8.6 \pm 2.5$			<sup>1</sup> VLADIMIRSK...08	SPEC	$40 \pi^-p \rightarrow K_S^0 K_S^0 n$
					$+m\pi^0$
$<80$	90	ALDE	87C	GAM2	$38 \pi^-p \rightarrow \eta' \eta n$
<sup>1</sup> $J^P C = 2^+ +$ . Systematic uncertainties not evaluated					

 $f_J(2220)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \pi\pi$	not seen
$\Gamma_2 \pi^+\pi^-$	not seen
$\Gamma_3 K\bar{K}$	not seen
$\Gamma_4 p\bar{p}$	not seen
$\Gamma_5 \gamma\gamma$	not seen
$\Gamma_6 \eta\eta'(958)$	seen
$\Gamma_7 \phi\phi$	not seen
$\Gamma_8 \eta\eta$	not seen

 $f_J(2220) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$ 

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.4$	95	<sup>1</sup> ACCIARRI	01H L3	$\gamma\gamma \rightarrow K_S^0 K_S^0, E_{\text{cm}}^{\text{ee}} = 91, 183-209 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 5.6$	95	<sup>1</sup> GODANG	97 CLE2	$\gamma\gamma \rightarrow K_S^0 K_S^0$
$< 86$	95	<sup>1</sup> ALBRECHT	90G ARG	$\gamma\gamma \rightarrow K^+K^-$
$< 1000$	95	<sup>2</sup> ALTHOFF	85B TASS	$\gamma\gamma, K\bar{K}\pi$
<b><math>\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}</math></b>				<b><math>\Gamma_1\Gamma_5/\Gamma</math></b>
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.5</math></b>	95	ALAM	98C CLE2	$\gamma\gamma \rightarrow \pi^+\pi^-$
<sup>1</sup> Assuming $J^P = 2^+$ . <sup>2</sup> True for $J^P = 0^+$ and $J^P = 2^+$ .				

 $f_J(2220) \Gamma(i)\Gamma(p\bar{p})/\Gamma^2(\text{total})$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 18</math></b>	95	<sup>1</sup> AMSLER	01 CBAR	$1.4-1.5 p\bar{p} \rightarrow \pi^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< (11-42)$	99	<sup>2</sup> HASAN	96 SPEC	$1.35-1.55 p\bar{p} \rightarrow \pi^+\pi^-$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 6</math></b>	95	<sup>3</sup> EVANGELIS...	98 SPEC	$1.1-2.0 p\bar{p} \rightarrow \phi\phi$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4</math></b>	95	<sup>1</sup> AMSLER	01 CBAR	$1.4-1.5 p\bar{p} \rightarrow \eta\eta$
<sup>1</sup> For $J^P = 2^+$ in the mass range 2222-2240 MeV and the total width between 10 and 20 MeV. <sup>2</sup> For $J^P = 2^+$ and $J^P = 4^+$ in the mass range 2220-2245 MeV and the total width of 15 MeV. <sup>3</sup> For $J^P = 2^+$ , the mass of 2235 MeV and the total width of 15 MeV.				

 $f_J(2220)$  BRANCHING RATIOS

VALUE	DOCUMENT ID	COMMENT
<b>not seen</b>	<sup>1</sup> DOBBS 15	$J/\psi \rightarrow \gamma\pi\pi$
not seen	<sup>1</sup> DOBBS 15	$\psi(25) \rightarrow \gamma K\pi$
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.		

VALUE	DOCUMENT ID	COMMENT
<b>not seen</b>	<sup>1</sup> DOBBS 15	$J/\psi \rightarrow \gamma K\bar{K}$
not seen	<sup>1</sup> DOBBS 15	$\psi(25) \rightarrow \gamma K\bar{K}$
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.		

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.0 \pm 0.5</math></b>	BAI	96B BES	$e^+e^- \rightarrow J/\psi \rightarrow \gamma 2\pi, K\bar{K}$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen		<sup>1</sup> AUBERT 07AV	BABR	$B \rightarrow p\bar{p}K^{(*)}$
not seen		WANG 05A	BELL	$B^+ \rightarrow \bar{p}p K^+$
$< 3.0$	95	<sup>2</sup> EVANGELIS...	97 SPEC	$1.96-2.40 p\bar{p} \rightarrow K_S^0 K_S^0$
$< 1.1$	99.7	<sup>3</sup> BARNES	93 SPEC	$1.3-1.57 p\bar{p} \rightarrow K_S^0 K_S^0$
$< 2.6$	99.7	<sup>3</sup> BARDIN	87 CNTR	$1.3-1.5 p\bar{p} \rightarrow K^+K^-$
$< 3.6$	99.7	<sup>3</sup> SCULLI	87 CNTR	$1.29-1.55 p\bar{p} \rightarrow K^+K^-$
<sup>1</sup> Assuming $\Gamma < 30 \text{ MeV}$ . <sup>2</sup> Assuming $\Gamma \sim 20 \text{ MeV}$ , $J^P = 2^+$ and $B(f_J(2220) \rightarrow K\bar{K}) = 100\%$ . <sup>3</sup> Assuming $\Gamma = 30-35 \text{ MeV}$ , $J^P = 2^+$ and $B(f_J(2220) \rightarrow K\bar{K}) = 100\%$ .				

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.17 \pm 0.09</math></b>	BAI	96B BES	$e^+e^- \rightarrow J/\psi \rightarrow \gamma p\bar{p}, K\bar{K}$

 $f_J(2220)$  REFERENCES

DOBBS 15	PR D91 052006	S. Dobbs <i>et al.</i>	(NWES)
VLADIMIRSK... 08	PAN 71 2129	V.V. Vladimirovsky <i>et al.</i>	(ITEP)
AUBERT 07AV	Translated from YAF 71 2166	B. Aubert <i>et al.</i>	(BABAR Collab.)
WANG 05A	PL B617 141	M.-Z. Wang <i>et al.</i>	(BELLE Collab.)
PDG 04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
ACCIARRI 01H	PL B501 173	M. Acciarri <i>et al.</i>	(L3 Collab.)
AMSLER 01	PL B520 175	C. Amisler <i>et al.</i>	(Crystal Barrel Collab.)
ALAM 98C	PRL 81 3328	M.S. Alam <i>et al.</i>	(CLEO Collab.)
BAI 98B	PRL 81 1179	J.Z. Bai <i>et al.</i>	(BES Collab.)
EVANGELIS... 98	PR D57 5370	C. Evangelista <i>et al.</i>	(JETSET Collab.)
EVANGELIS... 97	PR D56 3803	C. Evangelista <i>et al.</i>	(LEAR Collab.)
GODANG 97	PRL 79 3829	R. Godang <i>et al.</i>	(CLEO Collab.)
BAI 96B	PRL 76 3502	J.Z. Bai <i>et al.</i>	(BES Collab.)
HASAN 96	PL B388 376	A. Hasan, D.V. Bugg	(BRUN, LOQM)
BARNES 93	PL B309 469	P.D. Barnes <i>et al.</i>	(PS185 Collab.)
ALBRECHT 90G	ZPHY C48 183	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ASTON 88F	PL B215 199	D. Aston <i>et al.</i>	(SLAC, NAGO, CIN C, INUS)JP
BOLONKIN 88	NP B309 426	B.V. Bolonkin <i>et al.</i>	(ITEP, SERP)
ALDE 87C	SJNP 45 255	D. Alde <i>et al.</i>	
	Translated from YAF 45 405.		
BARDIN 87	PL B195 292	G. Bardin <i>et al.</i>	(SACL, FERR, CERN, PADO+)
SCULLI 87	PRL 58 1715	J. Sculli <i>et al.</i>	(NYU, BNL)
ALDE 86B	PL B177 120	D.M. Alde <i>et al.</i>	(SERP, BELG, LANL, LAPP)
BALTRUSAIT... 85D	PRL 56 107	R.M. Baltrusaitis	(CIT, UCSC, ILL, SLAC+)
ALTHOFF 85B	ZPHY C29 189	M. Althoff <i>et al.</i>	(TASSO Collab.)

## OTHER RELATED PAPERS

DEL-AMO-SA... 100	PRL 105 172001	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
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 $\eta(2225)$ 

$$I^G(J^{PC}) = 0^+(0^-+)$$

OMITTED FROM SUMMARY TABLE

Seen in  $J/\psi \rightarrow \gamma\phi\phi$ . Possibly seen in  $B \rightarrow \phi\phi K$  by LEES 11A. $\eta(2225)$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2226 \pm 16</math></b>		<b>OUR AVERAGE</b>		
$2240^{+30+30}_{-20-20}$	$196 \pm 19$	ABLIKIM	08i BES	$J/\psi \rightarrow \gamma K^+K^- K_S^0 K_L^0$
$2230 \pm 25 \pm 15$		BAI	90B MRK3	$J/\psi \rightarrow \gamma K^+K^- K^+K^-$
$2214 \pm 20 \pm 13$		BAI	90B MRK3	$J/\psi \rightarrow \gamma K^+K^- K_S^0 K_L^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 2220$		BISELLO	86B DM2	$J/\psi \rightarrow \gamma K^+K^- K^+K^-$

# Meson Particle Listings

## $\eta(2225)$ , $\rho_3(2250)$ , $f_2(2300)$

### $\eta(2225)$ WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>185^{+70}_{-40}</math></b>	<b>OUR AVERAGE</b>			
$190 \pm 30^{+60}_{-40}$	$196 \pm 19$	ABLIKIM	08i BES	$J/\psi \rightarrow \gamma K^+ K^- K_S^0 K_L^0$ (BES Collab.)
$150^{+300}_{-60} \pm 60$		BAI	90b MRK3	$J/\psi \rightarrow \gamma K^+ K^- K^+ K^-$ (Mark III Collab.)
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 80$		BISELLO	86b DM2	$J/\psi \rightarrow \gamma K^+ K^- K^+ K^-$ (DM2 Collab.)

### $\eta(2225)$ REFERENCES

LEES	11A	PR D84 012001	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABLIKIM	08i	PL B662 330	M. Ablikim <i>et al.</i>	(BES Collab.)
BAI	90b	PRL 65 1309	Z. Bai <i>et al.</i>	(Mark III Collab.)
BISELLO	86b	PL B179 294	D. Bisello <i>et al.</i>	(DM2 Collab.)

## $\rho_3(2250)$

$$I^G(J^{PC}) = 1^+(3^{--})$$

OMITTED FROM SUMMARY TABLE

Contains results mostly from formation experiments. For further production experiments see the Further States entry. See also  $\rho(2150)$ ,  $f_2(2150)$ ,  $f_4(2300)$ ,  $\rho_5(2350)$ .

### $\rho_3(2250)$ MASS

#### $\overline{p}p \rightarrow \pi\pi \text{ or } K\overline{K}$

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 2232$	HASAN	94	RVUE	$\overline{p}p \rightarrow \pi\pi$
$\sim 2090$	1 OAKDEN	94	RVUE	$0.36\text{--}1.55 \overline{p}p \rightarrow \pi\pi$
$\sim 2250$	2 MARTIN	80b	RVUE	
$\sim 2300$	2 MARTIN	80c	RVUE	
$\sim 2140$	3 CARTER	78b	CNTR 0	$0.7\text{--}2.4 \overline{p}p \rightarrow K^- K^+$
$\sim 2150$	4 CARTER	77	CNTR 0	$0.7\text{--}2.4 \overline{p}p \rightarrow \pi\pi$

- <sup>1</sup> See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J = 3$  to be important but not significantly resonant.  
<sup>2</sup>  $I(J^P) = 1(3^-)$  from simultaneous analysis of  $p\overline{p} \rightarrow \pi^-\pi^+$  and  $\pi^0\pi^0$ .  
<sup>3</sup>  $I = 0, 1$ .  $J^P = 3^-$  from Barrelet-zero analysis.  
<sup>4</sup>  $I(J^P) = 1(3^-)$  from amplitude analysis.

### S-CHANNEL $\overline{N}N$

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2260 \pm 20$	5 ANISOVICH	02	SPEC	$0.6\text{--}1.9 p\overline{p} \rightarrow \omega\pi^0$ , $\omega\eta\pi^0$ , $\pi^+\pi^-$
$\sim 2190$	6 CUTTS	78b	CNTR	$0.97\text{--}3 \overline{p}p \rightarrow \overline{N}N$
$2155 \pm 15$	6,7 COUPLAND	77	CNTR 0	$0.7\text{--}2.4 \overline{p}p \rightarrow \overline{p}p$
$2193 \pm 2$	6,8 ALSPECTOR	73	CNTR	$\overline{p}p$ S channel
$2190 \pm 10$	9 ABRAMS	70	CNTR	S channel $\overline{p}N$

- <sup>5</sup> From the combined analysis of ANISOVICH 00j, ANISOVICH 01d, ANISOVICH 01e, and ANISOVICH 02.  
<sup>6</sup> Isospins 0 and 1 not separated.  
<sup>7</sup> From a fit to the total elastic cross section.  
<sup>8</sup> Referred to as  $T$  or  $\overline{T}$  region by ALSPECTOR 73.  
<sup>9</sup> Seen as bump in  $I = 1$  state. See also COOPER 68. PEASLEE 75 confirm  $\overline{p}p$  results of ABRAMS 70, no narrow structure.

#### $\pi^-p \rightarrow \eta\pi\pi$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2290 \pm 20 \pm 30$	AMELIN	00 VES	$37 \pi^-p \rightarrow \eta\pi^+\pi^-n$

### $\rho_3(2250)$ WIDTH

#### $\overline{p}p \rightarrow \pi\pi \text{ or } K\overline{K}$

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 220$	HASAN	94	RVUE	$\overline{p}p \rightarrow \pi\pi$
$\sim 60$	10 OAKDEN	94	RVUE	$0.36\text{--}1.55 \overline{p}p \rightarrow \pi\pi$
$\sim 250$	11 MARTIN	80b	RVUE	
$\sim 200$	11 MARTIN	80c	RVUE	
$\sim 150$	12 CARTER	78b	CNTR 0	$0.7\text{--}2.4 \overline{p}p \rightarrow K^- K^+$
$\sim 200$	13 CARTER	77	CNTR 0	$0.7\text{--}2.4 \overline{p}p \rightarrow \pi\pi$

- <sup>10</sup> See however KLOET 96 who fit  $\pi^+\pi^-$  only and find waves only up to  $J = 3$  to be important but not significantly resonant.  
<sup>11</sup>  $I(J^P) = 1(3^-)$  from simultaneous analysis of  $p\overline{p} \rightarrow \pi^-\pi^+$  and  $\pi^0\pi^0$ .  
<sup>12</sup>  $I = 0, 1$ .  $J^P = 3^-$  from Barrelet-zero analysis.  
<sup>13</sup>  $I(J^P) = 1(3^-)$  from amplitude analysis.

### S-CHANNEL $\overline{N}N$

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$160 \pm 25$	14 ANISOVICH	02	SPEC	$0.6\text{--}1.9 p\overline{p} \rightarrow \omega\pi^0$ , $\omega\eta\pi^0$ , $\pi^+\pi^-$
$135 \pm 75$	15,16 COUPLAND	77	CNTR 0	$0.7\text{--}2.4 p\overline{p} \rightarrow \overline{p}p$
$98 \pm 8$	16 ALSPECTOR	73	CNTR	$\overline{p}p$ S channel
$\sim 85$	17 ABRAMS	70	CNTR	S channel $\overline{p}N$

- <sup>14</sup> From the combined analysis of ANISOVICH 00j, ANISOVICH 01d, ANISOVICH 01e, and ANISOVICH 02.  
<sup>15</sup> From a fit to the total elastic cross section.  
<sup>16</sup> Isospins 0 and 1 not separated.  
<sup>17</sup> Seen as bump in  $I = 1$  state. See also COOPER 68. PEASLEE 75 confirm  $\overline{p}p$  results of ABRAMS 70, no narrow structure.

#### $\pi^-p \rightarrow \eta\pi\pi$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$230 \pm 50 \pm 80$	AMELIN	00 VES	$37 \pi^-p \rightarrow \eta\pi^+\pi^-n$

### $\rho_3(2250)$ REFERENCES

ANISOVICH	02	PL B542 8	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01D	PL B508 6	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01E	PL B513 281	A.V. Anisovich <i>et al.</i>	
AMELIN	00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
KLOET	96	PR D53 6120	W.M. Kloet, F. Myhrer	(RUTG, NORD)
HASAN	94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)
OAKDEN	94	NP A574 731	M.N. Oakden, M.R. Pennington	(DURH)
MARTIN	80B	NP B176 355	B.R. Martin, D. Morgan	(LOUC, RHEL)JP
MARTIN	80C	NP B169 216	A.D. Martin, M.R. Pennington	(DURH)JP
CARTER	78B	NP B141 467	A.A. Carter	(LOQM)
CUTTS	78B	PR D17 16	D. Cutts <i>et al.</i>	(STON, WISC)
CARTER	77	PL 67B 117	A.A. Carter <i>et al.</i>	(LOQM, RHEL)JP
COUPLAND	77	PL 71B 460	M. Coupland <i>et al.</i>	(LOQM, RHEL)
PEASLEE	75	PL 57B 189	D.C. Peaslee <i>et al.</i>	(CANB, BARI, BROW+)
ALSPECTOR	73	PRL 30 511	J. Alspector <i>et al.</i>	(RUTG, UPNJ)
ABRAMS	70	PR D1 1917	R.J. Abrams <i>et al.</i>	(BNL)
COOPER	68	PRL 20 1059	W.A. Cooper <i>et al.</i>	(ANL)

## $f_2(2300)$

$$I^G(J^{PC}) = 0^+(2^{++})$$

### $f_2(2300)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>2297 \pm 28</math></b>	1 ETKIN	88	MPS $22 \pi^-p \rightarrow \phi\phi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2243^{+7+3}_{-6-29}$	UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$
$2270 \pm 12$	VLADIMIRSK..06	SPEC	$40 \pi^-p \rightarrow K_S^0 K_S^0 n$
$2327 \pm 9 \pm 6$	ABE	04	BELL $10.6 e^+e^- \rightarrow e^+e^- K^+ K^-$
$2231 \pm 10$	BOOTH	86	OMEG $85 \pi^- \text{Be} \rightarrow 2\phi \text{Be}$
$2220^{+90}_{-20}$	LINDENBAUM	84	RVUE
$2320 \pm 40$	ETKIN	82	MPS $22 \pi^-p \rightarrow 2\phi n$

- <sup>1</sup> Includes data of ETKIN 85. The percentage of the resonance going into  $\phi\phi 2^{++} + S_2$ ,  $D_2$ , and  $D_0$  is  $6^{+15}_{-5}$ ,  $25^{+18}_{-14}$ , and  $69^{+16}_{-27}$ , respectively.

### $f_2(2300)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>149 \pm 41</math></b>	2 ETKIN	88	MPS $22 \pi^-p \rightarrow \phi\phi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$145 \pm 12^{+27}_{-34}$	UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$
$90 \pm 29$	VLADIMIRSK..06	SPEC	$40 \pi^-p \rightarrow K_S^0 K_S^0 n$
$275 \pm 36 \pm 20$	ABE	04	BELL $10.6 e^+e^- \rightarrow e^+e^- K^+ K^-$
$133 \pm 50$	BOOTH	86	OMEG $85 \pi^- \text{Be} \rightarrow 2\phi \text{Be}$
$200 \pm 50$	LINDENBAUM	84	RVUE
$220 \pm 70$	ETKIN	82	MPS $22 \pi^-p \rightarrow 2\phi n$

- <sup>2</sup> Includes data of ETKIN 85.

### $f_2(2300)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \phi\phi$	seen
$\Gamma_2 K\overline{K}$	seen
$\Gamma_3 \gamma\gamma$	seen

See key on page 601

# Meson Particle Listings

## $f_2(2300)$ , $f_4(2300)$ , $f_0(2330)$

### $f_2(2300)$ $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2\Gamma_3/\Gamma$
VALUE (eV)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.2^{+0.5+1.3}_{-0.4-2.2}$	UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$	
$44 \pm 6 \pm 12$	<sup>3</sup> ABE	04	BELL $10.6 e^+ e^- \rightarrow e^+ e^- K^+ K^-$	
<sup>3</sup> Assuming spin 2.				

### $f_2(2300)$ REFERENCES

UEHARA 13	PTEP 2013 123C01	S. Uehara <i>et al.</i>	(BELLE Collab.)
VLADIMIRSK... 06	PAN 69 493	V.V. Vladimirovsky <i>et al.</i>	(ITEP, Moscow)
ABE 04	Translated from YAF 69 515	K. Abe <i>et al.</i>	(BELLE Collab.)
ETKIN 88	EPJ C32 323	A. Etkin <i>et al.</i>	(BNL, CUNY)
BOOTH 86	PL B273 677	P.S.L. Booth <i>et al.</i>	(LIVP, GLAS, CERN)
ETKIN 85	PL 165B 217	A. Etkin <i>et al.</i>	(BNL, CUNY)
LINDENBAUM 84	CNPP 13 285	S.J. Lindenbaum	(CUNY)
ETKIN 82	PRL 49 1620	A. Etkin <i>et al.</i>	(BNL, CUNY)

$f_4(2300)$

$$I^G(J^{PC}) = 0^+(4^{++})$$

OMITTED FROM SUMMARY TABLE

This entry was previously called  $U_0(2350)$ . Contains results mostly from formation experiments. For further production experiments see the Further States entry. See also  $\rho(2150)$ ,  $f_2(2150)$ ,  $\rho_3(2250)$ ,  $\rho_5(2350)$ .

### $f_4(2300)$ MASS

### $\bar{p}p \rightarrow \pi\pi\alpha\bar{K}K$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$\sim 2314$	HASAN 94	RVUE	$\bar{p}p \rightarrow \pi\pi$
$\sim 2300$	<sup>1</sup> MARTIN 80B	RVUE	
$\sim 2300$	<sup>1</sup> MARTIN 80C	RVUE	
$\sim 2340$	<sup>2</sup> CARTER 78B	CNTR	$0.7-2.4 \bar{p}p \rightarrow K^- K^+$
$\sim 2330$	DULUDE 78B	OSPK	$1-2 \bar{p}p \rightarrow \pi^0 \pi^0$
$\sim 2310$	<sup>3</sup> CARTER 77	CNTR	$0.7-2.4 \bar{p}p \rightarrow \pi\pi$
<sup>1</sup> $I(J^P) = 0(4^+)$ from simultaneous analysis of $p\bar{p} \rightarrow \pi^- \pi^+$ and $\pi^0 \pi^0$ .			
<sup>2</sup> $I(J^P) = 0(4^+)$ from Barrelet-zero analysis.			
<sup>3</sup> $I(J^P) = 0(4^+)$ from amplitude analysis.			

### S-CHANNEL $\bar{p}p \alpha \bar{N}N$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2283 \pm 17$	<sup>4</sup> ANISOVICH 00J	SPEC	
$\sim 2380$	<sup>5</sup> CUTTS 78B	CNTR	$0.97-3 \bar{p}p \rightarrow \bar{N}N$
$2345 \pm 15$	<sup>5,6</sup> COUPLAND 77	CNTR	$0.7-2.4 \bar{p}p \rightarrow \bar{p}p$
$2359 \pm 2$	<sup>5,7</sup> ALSPECTOR 73	CNTR	$\bar{p}p$ S channel
$2375 \pm 10$	ABRAMS 70	CNTR	S channel $\bar{N}N$
<sup>4</sup> From the combined analysis of ANISOVICH 99c and ANISOVICH 99f on $\bar{p}p \rightarrow \eta\pi^0\pi^0$ , $\pi^0\pi^0$ , $\eta\eta$ , $\eta\eta'$ , $\pi^+\pi^-$ .			
<sup>5</sup> Isospins 0 and 1 not separated.			
<sup>6</sup> From a fit to the total elastic cross section.			
<sup>7</sup> Referred to as $U$ or $U$ region by ALSPECTOR 73.			

### $\pi^- p \rightarrow \eta\pi\pi n$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2330 \pm 20 \pm 40$	AMELIN 00	VES	$37 \pi^- p \rightarrow \eta\pi^+\pi^- n$

### $pp$ CENTRAL PRODUCTION

VALUE (MeV)	DOCUMENT ID	COMMENT
<b>2320 <math>\pm</math> 60 OUR ESTIMATE</b>		
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$2332 \pm 15$	BARBERIS 00f	$450 pp \rightarrow p_f \omega p_S$

### $f_4(2300)$ WIDTH

### $\bar{p}p \rightarrow \pi\pi\alpha\bar{K}K$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$\sim 278$	HASAN 94	RVUE	$\bar{p}p \rightarrow \pi\pi$
$\sim 200$	<sup>8</sup> MARTIN 80C	RVUE	
$\sim 150$	<sup>9</sup> CARTER 78B	CNTR	$0.7-2.4 \bar{p}p \rightarrow K^- K^+$
$\sim 210$	<sup>10</sup> CARTER 77	CNTR	$0.7-2.4 \bar{p}p \rightarrow \pi\pi$
<sup>8</sup> $I(J^P) = 0(4^+)$ from simultaneous analysis of $p\bar{p} \rightarrow \pi^- \pi^+$ and $\pi^0 \pi^0$ .			
<sup>9</sup> $I(J^P) = 0(4^+)$ from Barrelet-zero analysis.			
<sup>10</sup> $I(J^P) = 0(4^+)$ from amplitude analysis.			

### S-CHANNEL $\bar{p}p \alpha \bar{N}N$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$310 \pm 25$	<sup>11</sup> ANISOVICH 00J	SPEC	
$135^{+150}_{-65}$	<sup>12,13</sup> COUPLAND 77	CNTR	$0.7-2.4 \bar{p}p \rightarrow \bar{p}p$
$165^{+18}_{-8}$	<sup>13</sup> ALSPECTOR 73	CNTR	$\bar{p}p$ S channel
$\sim 190$	ABRAMS 70	CNTR	S channel $\bar{N}N$
<sup>11</sup> From the combined analysis of ANISOVICH 99c and ANISOVICH 99f on $\bar{p}p \rightarrow \eta\pi^0\pi^0$ , $\pi^0\pi^0$ , $\eta\eta$ , $\eta\eta'$ , $\pi^+\pi^-$ .			
<sup>12</sup> From a fit to the total elastic cross section.			
<sup>13</sup> Isospins 0 and 1 not separated.			

### $\pi^- p \rightarrow \eta\pi\pi n$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$235 \pm 50 \pm 40$	AMELIN 00	VES	$37 \pi^- p \rightarrow \eta\pi^+\pi^- n$

### $pp$ CENTRAL PRODUCTION

VALUE (MeV)	DOCUMENT ID	COMMENT
<b>250 <math>\pm</math> 80 OUR ESTIMATE</b>		
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$260 \pm 57$	BARBERIS 00f	$450 pp \rightarrow p_f \omega p_S$

### $f_4(2300)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\rho\rho$	seen
$\Gamma_2$ $\omega\omega$	seen
$\Gamma_3$ $\eta\pi\pi$	seen
$\Gamma_4$ $\pi\pi$	seen
$\Gamma_5$ $K\bar{K}$	seen
$\Gamma_6$ $N\bar{N}$	seen

### $f_4(2300)$ BRANCHING RATIOS

$\Gamma(\rho\rho)/\Gamma(\omega\omega)$	DOCUMENT ID	COMMENT	$\Gamma_1/\Gamma_2$
VALUE			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.8 \pm 0.5$	BARBERIS 00f	$450 pp \rightarrow p_f \omega p_S$	

### $f_4(2300)$ REFERENCES

AMELIN 00	NP A668 83	D. Amelin <i>et al.</i>	(VES Collab.)
ANISOVICH 00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
BARBERIS 00f	PL B484 198	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ANISOVICH 99C	PL B452 173	A.V. Anisovich <i>et al.</i>	
ANISOVICH 99F	NP A551 253	A.V. Anisovich <i>et al.</i>	
HASAN 94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)
MARTIN 80B	NP B176 355	B.R. Martin, D. Morgan	(LOUC, RHEL)JP
MARTIN 80C	NP B169 216	A.D. Martin, M.R. Pennington	(DURH)JP
CARTER 78B	NP B141 467	A.A. Carter	(LOQM)
CUTTS 78B	PR D17 16	D. Cutts <i>et al.</i>	(STON, WISC)
DULUDE 78B	PL 79B 335	R.S. Dulude <i>et al.</i>	(BROW, MIT, BARI)JP
CARTER 77	PL 67B 117	A.A. Carter <i>et al.</i>	(LOQM, RHEL)JP
COUPLAND 77	PL 71B 460	M. Coupland <i>et al.</i>	(LOQM, RHEL)
ALSPECTOR 73	PRL 30 511	J. Alspector <i>et al.</i>	(RUTG, UPNJ)
ABRAMS 70	PR D1 1917	R.J. Abrams <i>et al.</i>	(BNL)

$f_0(2330)$

$$I^G(J^{PC}) = 0^+(0^{++})$$

OMITTED FROM SUMMARY TABLE

### $f_0(2330)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2314 \pm 25$	<sup>1</sup> BUGG 04A	RVUE	
$2337 \pm 14$	ANISOVICH 00J	SPEC	$2.0 \bar{p}p \rightarrow \pi\pi, \eta\eta$
$\sim 2321$	HASAN 94	RVUE	$\bar{p}p \rightarrow \pi\pi$
<sup>1</sup> Partial wave analysis of the data on $p\bar{p} \rightarrow \bar{\Lambda}\Lambda$ from BARNES 00.			

### $f_0(2330)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$144 \pm 20$	<sup>2</sup> BUGG 04A	RVUE	
$217 \pm 33$	ANISOVICH 00J	SPEC	$2.0 \bar{p}p \rightarrow \pi\pi, \eta\eta$
$\sim 223$	HASAN 94	RVUE	$\bar{p}p \rightarrow \pi\pi$
<sup>2</sup> Partial wave analysis of the data on $p\bar{p} \rightarrow \bar{\Lambda}\Lambda$ from BARNES 00.			

### $f_0(2330)$ REFERENCES

BUGG 04A	EPJ C36 161	D.V. Bugg	
ANISOVICH 00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
BARNES 00	PR C62 055203	P.D. Barnes <i>et al.</i>	
HASAN 94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)

# Meson Particle Listings

## $f_2(2340)$ , $\rho_5(2350)$

$f_2(2340)$				
$J^G(J^{PC}) = 0^+(2^{++})$				
$f_2(2340)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2345<math>\pm</math>50<sub>-40</sub> OUR AVERAGE</b>				
2362 $\pm$ 31 <sub>-30</sub> <sup>+140</sup> <sub>-63</sub>	5.5k	<sup>1</sup> ABLIKIM	13N BES3	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\eta\eta$
2339 $\pm$ 55		<sup>2</sup> ETKIN	88 MPS	$22\pi^-p \rightarrow \phi\phi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2350 $\pm$ 7	80k	<sup>3</sup> UMAN	06 E835	$5.2\bar{p}p \rightarrow \eta\eta\pi^0$
2392 $\pm$ 10		BOOTH	86 OMEG	$85\pi^-Be \rightarrow 2\phi Be$
2360 $\pm$ 20		LINDENBAUM	84 RVUE	
<sup>1</sup> From partial wave analysis including all possible combinations of $0^{++}$ , $2^{++}$ , and $4^{++}$ resonances.				
<sup>2</sup> Includes data of ETKIN 85. The percentage of the resonance going into $\phi\phi\ 2^{++}\ S_2$ , $D_2$ , and $D_0$ is $37\pm 19$ , $4^{+12}_{-4}$ , and $59^{+21}_{-19}$ , respectively.				
<sup>3</sup> Statistical error only.				

$f_2(2340)$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>322<math>\pm</math>70<sub>-60</sub> OUR AVERAGE</b>				
334 $\pm$ 62 <sub>-54</sub> <sup>+165</sup> <sub>-100</sub>	5.5k	<sup>4</sup> ABLIKIM	13N BES3	$e^+e^- \rightarrow J/\psi \rightarrow \gamma\eta\eta$
319 $\pm$ 81 <sub>-69</sub>		<sup>5</sup> ETKIN	88 MPS	$22\pi^-p \rightarrow \phi\phi n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
218 $\pm$ 16	80k	<sup>6</sup> UMAN	06 E835	$5.2\bar{p}p \rightarrow \eta\eta\pi^0$
198 $\pm$ 50		BOOTH	86 OMEG	$85\pi^-Be \rightarrow 2\phi Be$
150 $\pm$ 150 <sub>-50</sub>		LINDENBAUM	84 RVUE	
<sup>4</sup> From partial wave analysis including all possible combinations of $0^{++}$ , $2^{++}$ , and $4^{++}$ resonances.				
<sup>5</sup> Includes data of ETKIN 85.				
<sup>6</sup> Statistical error only.				

$f_2(2340)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $\phi\phi$	seen	
$\Gamma_2$ $\eta\eta$	seen	

$f_2(2340)$ BRANCHING RATIOS				
$\Gamma(\eta\eta)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	UMAN	06 E835	$5.2\bar{p}p \rightarrow \eta\eta\pi^0$	

$f_2(2340)$ REFERENCES				
ABLIKIM	13N	PR D87 092009	Ablikim M. <i>et al.</i>	(BES III Collab.)
UMAN	06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
ETKIN	88	PL B201 568	A. Etkin <i>et al.</i>	(BNL, CUNY)
BOOTH	86	NP B273 477	P.S.L. Booth <i>et al.</i>	(LIVP, GLAS, CERN)
ETKIN	85	PL 165B 217	A. Etkin <i>et al.</i>	(BNL, CUNY)
LINDENBAUM	84	CNPP 13 285	S.J. Lindenbaum	(CUNY)

$\rho_5(2350)$		$J^G(J^{PC}) = 1^+(5^{--})$
OMITTED FROM SUMMARY TABLE		
This entry was previously called $U_1(2400)$ . See also $\rho(2150)$ , $f_2(2150)$ , $\rho_3(2250)$ , $f_4(2300)$ .		

$\rho_5(2350)$ MASS				
$\pi^-p \rightarrow \omega\pi^0 n$				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>2330<math>\pm</math>35</b>	ALDE	95	GAM2	$38\pi^-p \rightarrow \omega\pi^0 n$
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 2303$	HASAN	94	RVUE	$\bar{p}p \rightarrow \pi\pi$
$\sim 2300$	<sup>1</sup> MARTIN	80B	RVUE	
$\sim 2250$	<sup>1</sup> MARTIN	80C	RVUE	
$\sim 2500$	<sup>2</sup> CARTER	78B	CNTR 0	$0.7\text{--}2.4\bar{p}p \rightarrow K^-K^+$
$\sim 2480$	<sup>3</sup> CARTER	77	CNTR 0	$0.7\text{--}2.4\bar{p}p \rightarrow \pi\pi$

<b>S-CHANNEL <math>\bar{N}N</math></b>				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2300 $\pm$ 45	<sup>4</sup> ANISOVICH	02	SPEC	$0.6\text{--}1.9\ p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$
2295 $\pm$ 30	ANISOVICH	00J	SPEC	
$\sim 2380$	<sup>5</sup> CUTTS	78B	CNTR	$0.97\text{--}3\ \bar{p}p \rightarrow \bar{N}N$
2345 $\pm$ 15	<sup>5,6</sup> COUPLAND	77	CNTR 0	$0.7\text{--}2.4\ \bar{p}p \rightarrow \bar{p}p$
2359 $\pm$ 2	<sup>5,7</sup> ALSPECTOR	73	CNTR	$\bar{p}p$ S channel
2350 $\pm$ 10	<sup>8</sup> ABRAMS	70	CNTR	S channel $\bar{N}N$
2360 $\pm$ 25	<sup>9</sup> OH	70B	HDBC $-0$	$\bar{p}(pn), K^*K\ 2\pi$

$\pi^-p \rightarrow K^+K^-n$				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2307 $\pm$ 6	ALPER	80	CNTR 0	$62\pi^-p \rightarrow K^+K^-n$
<sup>1</sup> $I(J^P) = 1(5^-)$ from simultaneous analysis of $p\bar{p} \rightarrow \pi^-\pi^+$ and $\pi^0\pi^0$ .				
<sup>2</sup> $I = 0(1)$ ; $J^P = 5^-$ from Barrelet-zero analysis.				
<sup>3</sup> $I(J^P) = 1(5^-)$ from amplitude analysis.				
<sup>4</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.				
<sup>5</sup> Isospins 0 and 1 not separated.				
<sup>6</sup> From a fit to the total elastic cross section.				
<sup>7</sup> Referred to as $U$ or $U$ region by ALSPECTOR 73.				
<sup>8</sup> For $I = 1\ \bar{N}N$ .				
<sup>9</sup> No evidence for this bump seen in the $\bar{p}p$ data of CHAPMAN 71B. Narrow state not confirmed by OH 73 with more data.				

$\rho_5(2350)$ WIDTH				
$\pi^-p \rightarrow \omega\pi^0 n$				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>400<math>\pm</math>100</b>	ALDE	95	GAM2	$38\pi^-p \rightarrow \omega\pi^0 n$
$\bar{p}p \rightarrow \pi\pi\ \alpha\ \bar{K}K$				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 169$	HASAN	94	RVUE	$\bar{p}p \rightarrow \pi\pi$
$\sim 250$	<sup>10</sup> MARTIN	80B	RVUE	
$\sim 300$	<sup>10</sup> MARTIN	80C	RVUE	
$\sim 150$	<sup>11</sup> CARTER	78B	CNTR 0	$0.7\text{--}2.4\ \bar{p}p \rightarrow K^-K^+$
$\sim 210$	<sup>12</sup> CARTER	77	CNTR 0	$0.7\text{--}2.4\ \bar{p}p \rightarrow \pi\pi$

<b>S-CHANNEL <math>\bar{N}N</math></b>				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
260 $\pm$ 75	<sup>13</sup> ANISOVICH	02	SPEC	$0.6\text{--}1.9\ p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$
235 $\pm$ 65 <sub>-40</sub>	ANISOVICH	00J	SPEC	
135 $\pm$ 150 <sub>-65</sub>	<sup>14,15</sup> COUPLAND	77	CNTR 0	$0.7\text{--}2.4\ \bar{p}p \rightarrow \bar{p}p$
165 $\pm$ 18 <sub>-8</sub>	<sup>15</sup> ALSPECTOR	73	CNTR	$\bar{p}p$ S channel
$< 60$	<sup>16</sup> OH	70B	HDBC $-0$	$\bar{p}(pn), K^*K\ 2\pi$
$\sim 140$	ABRAMS	67C	CNTR	S channel $\bar{p}N$

$\pi^-p \rightarrow K^+K^-n$				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
245 $\pm$ 20	ALPER	80	CNTR 0	$62\pi^-p \rightarrow K^+K^-n$
<sup>10</sup> $I(J^P) = 1(5^-)$ from simultaneous analysis of $p\bar{p} \rightarrow \pi^-\pi^+$ and $\pi^0\pi^0$ .				
<sup>11</sup> $I = 0(1)$ ; $J^P = 5^-$ from Barrelet-zero analysis.				
<sup>12</sup> $I(J^P) = 1(5^-)$ from amplitude analysis.				
<sup>13</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.				
<sup>14</sup> From a fit to the total elastic cross section.				
<sup>15</sup> Isospins 0 and 1 not separated.				
<sup>16</sup> No evidence for this bump seen in the $\bar{p}p$ data of CHAPMAN 71B. Narrow state not confirmed by OH 73 with more data.				

$\rho_5(2350)$ REFERENCES				
ANISOVICH	02	PL B542 8	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01D	PL B508 6	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01E	PL B513 281	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
ALDE	95	ZPHY C66 379	D.M. Alde <i>et al.</i>	(GAMS Collab.)JP
HASAN	94	PL B334 215	A. Hasan, D.V. Bugg	(LOGM)
ALPER	80	PL 94B 422	B. Alper <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
MARTIN	80B	NP B176 355	B.R. Martin, D. Morgan	(LOUC, RHEL)JP
MARTIN	80C	NP B169 216	A.D. Martin, M.R. Pennington	(BURH)JP
CARTER	78B	NP B141 467	A.A. Carter	(LOGM)
CUTTS	78B	PR D17 16	D. Cutts <i>et al.</i>	(STON, WISC)
CARTER	77	PL 67B 117	A.A. Carter <i>et al.</i>	(LOQM, RHEL)JP
COUPLAND	77	PL 71B 460	M. Coupland <i>et al.</i>	(LOQM, RHEL)
ALSPECTOR	73	PRL 30 511	J. Alspector <i>et al.</i>	(RUTG, UPNJ)
OH	73	NP B51 57	B.Y. Oh <i>et al.</i>	(MSU)
CHAPMAN	71B	PR D1 1275	J.W. Chapman <i>et al.</i>	(MICH)
ABRAMS	70	PR D1 1917	R.J. Abrams <i>et al.</i>	(BNL)
OH	70B	PRL 24 1257	B.Y. Oh <i>et al.</i>	(MSU)
ABRAMS	67C	PRL 18 1209	R.J. Abrams <i>et al.</i>	(BNL)

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Meson Particle Listings

$a_6(2450)$ ,  $f_6(2510)$

$a_6(2450)$

$I^G(J^{PC}) = 1^-(6^{++})$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

$a_6(2450)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
$2450 \pm 130$	<sup>1</sup> CLELAND	82B	SPEC	$\pm$ 50 $\pi p \rightarrow K_S^0 K^\pm p$

<sup>1</sup> From an amplitude analysis.

$a_6(2450)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
$400 \pm 250$	<sup>2</sup> CLELAND	82B	SPEC	$\pm$ 50 $\pi p \rightarrow K_S^0 K^\pm p$

<sup>2</sup> From an amplitude analysis.

$a_6(2450)$  DECAY MODES

Mode
$\Gamma_1 \quad K \bar{K}$

$a_6(2450)$  REFERENCES

CLELAND	82B	NP B208 228	W.E. Cleland <i>et al.</i>	(DURH, GEVA, LAUS+)
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$f_6(2510)$

$I^G(J^{PC}) = 0^+(6^{++})$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

$f_6(2510)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$2469 \pm 29$ OUR AVERAGE	Error includes scale factor of 1.5. See the ideogram below.		
$2485 \pm 40$	<sup>1</sup> ANISOVICH	00J	SPEC 1.92–2.41 $p\bar{p}$
$2420 \pm 30$	ALDE	98	GAM4 100 $\pi^- p \rightarrow \pi^0 \pi^0 n$
$2510 \pm 30$	BINON	84B	GAM2 38 $\pi^- p \rightarrow n 2\pi^0$

<sup>1</sup> From the combined analysis of ANISOVICH 99c, ANISOVICH 99f, ANISOVICH 99j, ANISOVICH 99k, and ANISOVICH 00b.

			$\chi^2$
ANISOVICH	00J	SPEC	0.2
ALDE	98	GAM4	2.7
BINON	84B	GAM2	1.8
			4.7

(Confidence Level = 0.096)

$f_6(2510)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$283 \pm 40$ OUR AVERAGE	Error includes scale factor of 1.1.		
$410 \pm 90$	<sup>2</sup> ANISOVICH	00J	SPEC 1.92–2.41 $p\bar{p}$
$270 \pm 60$	ALDE	98	GAM4 100 $\pi^- p \rightarrow \pi^0 \pi^0 n$
$240 \pm 60$	BINON	84B	GAM2 38 $\pi^- p \rightarrow n 2\pi^0$

<sup>2</sup> From the combined analysis of ANISOVICH 99c, ANISOVICH 99f, ANISOVICH 99j, ANISOVICH 99k, and ANISOVICH 00b.

$f_6(2510)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \pi \pi$	$(6.0 \pm 1.0) \%$

$f_6(2510)$  BRANCHING RATIOS

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
$0.06 \pm 0.01$	<sup>3</sup> BINON	83c	GAM2 38 $\pi^- p \rightarrow n 4\gamma$	

<sup>3</sup> Assuming one pion exchange and using data of BOLOTOV 74.

$f_6(2510)$  REFERENCES

ANISOVICH	00B	NP A662 319	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99C	PL B452 173	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99F	NP A651 253	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99J	PL B471 271	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99K	PL B468 309	A.V. Anisovich <i>et al.</i>	
ALDE	98	EPJ A3 361	D. Alde <i>et al.</i>	(GAM4 Collab.)
Also		PAN 62 405	D. Alde <i>et al.</i>	(GAMS Collab.)
		Translated from YAF 62 446.		
BINON	84B	LNC 39 41	F.G. Binon <i>et al.</i>	(SERP, BELG, LAPP)JP
BINON	83C	SJNP 38 723	F.G. Binon <i>et al.</i>	(SERP, BRUX+)
		Translated from YAF 38 1199.		
BOLOTOV	74	PL 52B 489	V.N. Bolotov <i>et al.</i>	(SERP)

# Meson Particle Listings

## Further States

### OTHER LIGHT MESONS

#### Further States

OMITTED FROM SUMMARY TABLE

This section contains states observed by a single group or states poorly established that thus need confirmation.

QUANTUM NUMBERS, MASSES, WIDTHS, AND BRANCHING RATIOS

$X(360) \quad I^G(J^{PC}) = ?^?(?^{?+})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
$360 \pm 7 \pm 9$	$64 \pm 18$	2.3k	<sup>1</sup> ABRAAMYAN 09	CNTR	$2.75 \, d \, C \rightarrow \gamma \gamma X$
<sup>1</sup> Not seen in $p \, C \rightarrow \gamma \gamma X$ at 5.5 GeV/c.					

$X(1070) \quad I^G(J^{PC}) = ?^?(0^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1072 \pm 1$	$3.5 \pm 0.5$		<sup>2</sup> VLADIMIRSK...08		$40 \, \pi^- p \rightarrow K_S^0 K_S^0 n + m \pi^0$
<sup>2</sup> Supersedes GRIGOR'EV 05.					

$X(1110) \quad I^G(J^{PC}) = 0^+(\text{even}^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1107 \pm 4$	$111 \pm 8 \pm 15$		DAFTARI 87	DBC	$0. \, \overline{p} n \rightarrow \rho^- \pi^+ \pi^-$

$\eta(1200-1600) \quad I^G(J^{PC}) = 0^+(0^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1323 \pm 8$	$237 \pm 20$		VLADIMIRSK...06	SPEC	$40 \, \pi^- p \rightarrow K_S^0 K_S^0 n$
$1480^{+100}_{-150}$	$1030^{+80}_{-170}$		<sup>3</sup> ANISOVICH 03	SPEC	
$1530^{+90}_{-250}$	$560 \pm 40$		<sup>4</sup> ANISOVICH 03	SPEC	
<sup>3</sup> K-matrix pole from combined analysis of $\pi^- p \rightarrow \pi^0 \pi^0 n$ , $\pi^- p \rightarrow \pi^0 K \overline{K} n$ , $\pi^+ \pi^- \rightarrow \pi^+ \pi^-$ , $\overline{p} p \rightarrow \pi^0 \pi^0 \pi^0$ , $\pi^0 \eta \eta$ , $\pi^0 \pi^0 \eta$ , $\pi^+ \pi^- \pi^0$ , $K^+ K^- \pi^0$ , $K_S^0 K_S^0 \pi^0$ , $K^+ K_S^0 \pi^-$ at rest, $\overline{p} n \rightarrow \pi^- \pi^- \pi^+$ , $K_S^0 K^- \pi^0$ , $K_S^0 K_S^0 \pi^-$ at rest.					
<sup>4</sup> K-matrix pole from combined analysis of $\pi^- p \rightarrow \pi^0 \pi^0 n$ , $\pi^- p \rightarrow K \overline{K} n$ , $\overline{p} p \rightarrow \pi^0 \pi^0 \pi^0$ , $\pi^0 \eta \eta$ , $\pi^0 \pi^0 \eta$ at rest.					

$X(1420) \quad I^G(J^{PC}) = 2^+(0^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1420 \pm 20$	$160 \pm 10$		FILIPPI 00	OBLX	$0 \, \overline{p} p \rightarrow \pi^+ \pi^+ \pi^-$

$X(1545) \quad I^G(J^{PC}) = ?^?(?^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	COMMENT	
$1545 \pm 3$	$6.0 \pm 2.5$		<sup>5</sup> VLADIMIRSK...08	$40 \, \pi^- p \rightarrow K_S^0 K_S^0 n + m \pi^0$	
<sup>5</sup> Supersedes VLADIMIRSKII 00.					

$X(1575) \quad I^G(J^{PC}) = ?^?(1^{--})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1576^{+49+98}_{-55-91}$	$818^{+22+64}_{-23-133}$		<sup>6</sup> ABLIKIM 06s	BES	$J/\psi \rightarrow K^+ K^- \pi^0$
<sup>6</sup> A broad peak observed at $K^+ K^-$ invariant mass. Mass and width above are its pole position. The observed branching ratio is $B(J/\psi \rightarrow X \pi^0) B(X \rightarrow K^+ K^-) = (8.5 \pm 0.6^{+2.7}_{-3.6}) \times 10^{-4}$ .					

$X(1600) \quad I^G(J^{PC}) = 2^+(2^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1600 \pm 100$	$400 \pm 200$		<sup>7</sup> ALBRECHT 91F	ARG	$10.2 \, e^+ e^- \rightarrow e^+ e^- 2(\pi^+ \pi^-)$
<sup>7</sup> Our estimate.					

$X(1650) \quad I^G(J^{PC}) = 0^-(?^{? -})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
$1652 \pm 7$	$< 50$	100	PROKOSHKIN 96	GAM2	$32, 38 \, \pi p \rightarrow \omega \eta n$

$X(1730) \quad I^G(J^{PC}) = ?^?(?^{?+})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
$1731.0 \pm 1.2 \pm 2.0$	$3.2 \pm 0.8 \pm 1.3$	58	VLADIMIRSK...07	SPEC	$40 \, \pi^- p \rightarrow K_S^0 K_S^0 X$

$X(1750) \quad I^G(J^{PC}) = ?^?(1^{--})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1753.5 \pm 1.5 \pm 2.3$	$122.2 \pm 6.2 \pm 8.0$		LINK 02k	FOCS	$20-160 \, \gamma p \rightarrow K^+ K^- p$

$B(X(1750) \rightarrow \overline{K}^*(892)^0 K^0 \rightarrow K^\pm \pi^\mp K_S^0)/B(X(1750) \rightarrow K^+ K^-)$					
VALUE	CL%	DOCUMENT ID	TECN		
$< 0.065$	90	LINK	02k	FOCS	

$B(X(1750) \rightarrow \overline{K}^*(892)^\pm K^\mp \rightarrow K^\pm \pi^\mp K_S^0)/B(X(1750) \rightarrow K^+ K^-)$					
VALUE	CL%	DOCUMENT ID	TECN		
$< 0.183$	90	LINK	02k	FOCS	

$\rho_2(1750) \quad I^G(J^{PC}) = 0^+(2^{++})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
$1755 \pm 10$	$67 \pm 12$	870	<sup>8</sup> SCHEGELSKY 06A	RVUE	$\gamma \gamma \rightarrow K_S^0 K_S^0$

$\Gamma(K \overline{K})$					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT	
$17 \pm 5$	870	<sup>9</sup> SCHEGELSKY 06A	RVUE	$\gamma \gamma \rightarrow K_S^0 K_S^0$	

$\Gamma(\gamma \gamma)$					
VALUE (keV)	EVTs	DOCUMENT ID	TECN	COMMENT	
$0.13 \pm 0.04$	870	<sup>9</sup> SCHEGELSKY 06A	RVUE	$\gamma \gamma \rightarrow K_S^0 K_S^0$	

$\Gamma(\pi \pi)$					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT	
$1.3 \pm 1.0$	870	<sup>9</sup> SCHEGELSKY 06A	RVUE	$\gamma \gamma \rightarrow K_S^0 K_S^0$	

$\Gamma(\eta \eta)$					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT	
$2.0 \pm 0.5$	870	<sup>9</sup> SCHEGELSKY 06A	RVUE	$\gamma \gamma \rightarrow K_S^0 K_S^0$	
<sup>8</sup> From analysis of L3 data at 91 and 183-209 GeV.					
<sup>9</sup> From analysis of L3 data at 91 and 183-209 GeV and using SU(3) relations.					

$X(1775) \quad I^G(J^{PC}) = 1^-(?^{-+})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1763 \pm 20$	$192 \pm 60$		CONDO 91	SHF	$\gamma p \rightarrow (p \pi^+) (\pi^+ \pi^- \pi^-)$
$1787 \pm 18$	$118 \pm 60$		CONDO 91	SHF	$\gamma p \rightarrow n \pi^+ \pi^+ \pi^-$

$\eta(1800) \quad I^G(J^{PC}) = 0^+(0^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1795 \pm 7^{+23}_{-20}$	$95 \pm 10^{+78}_{-82}$		ABLIKIM 13J	BES3	$J/\psi \rightarrow \gamma \omega \phi$
$1812^{+19}_{-26} \pm 18$	$105 \pm 20 \pm 28$		<sup>10</sup> ABLIKIM 06J	BES2	$J/\psi \rightarrow \gamma \omega \phi$
<sup>10</sup> Not seen by LIU 09 in $B^\pm \rightarrow K^\pm \omega \phi$ .					

$X(1850 - 3100) \quad I^G(J^{PC}) = ?^?(1^{--})$					
$\Gamma(e^+ e^-) B(X \rightarrow \text{hadrons})$ (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
$< 120$	90	<sup>11</sup> ANASHIN 11	KEDR	$e^+ e^- \rightarrow \text{hadrons}$	
<sup>11</sup> This limit is center-of-mass energy dependent. We quote the most stringent one.					

$X(1855) \quad I^G(J^{PC}) = ?^?(?^{??})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1856.6 \pm 5$	$20 \pm 5$		BRIDGES 86D	SPEC	$0. \, \overline{p} d \rightarrow \pi \pi N$

$X(1870) \quad I^G(J^{PC}) = ?^?(2^{??})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1870 \pm 40$	$250 \pm 30$		ALDE 86D	GAM4	$100 \, \pi^- p \rightarrow 2 \eta X$

$a_3(1875) \quad I^G(J^{PC}) = 1^-(3^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1874 \pm 43 \pm 96$	$385 \pm 121 \pm 114$		CHUNG 02	B852	$18.3 \, \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$

$B(a_3(1875) \rightarrow f_2(1270) \pi)/B(a_3(1875) \rightarrow \rho \pi)$				
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.8 \pm 0.2$	<sup>12</sup> CHUNG 02	B852	$18.3 \, \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$	
<sup>12</sup> Using the observable fractions of 50.0% $\rho \pi$ , 56.5% $f_2 \pi$ , and 11.8% $\rho_3 \pi$ .				

$B(a_3(1875) \rightarrow \rho_3(1690)\pi)/B(a_3(1875) \rightarrow \rho\pi)$				
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.9 \pm 0.3$	<sup>13</sup> CHUNG	02	B852	$18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
<sup>13</sup> Using the observable fractions of 50.0% $\rho\pi$ , 56.5% $f_2\pi$ , and 11.8% $\rho_3\pi$ .				

$a_1(1930) \quad I^G(J^{PC}) = 1^-(1^{++})$					
MASS (MeV)	WIDTH (MeV)		DOCUMENT ID	TECN	COMMENT
$1930^{+30}_{-70}$	$155 \pm 45$		ANISOVICH 01F	SPEC	$2.0 \, \overline{p} p \rightarrow 3 \pi^0, \pi^0 \eta, \pi^0 \eta'$

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# Meson Particle Listings

## Further States

<b>X(1935)</b> $I^G(J^{PC}) = 1^+(1^{-?})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1935 ± 20	215 ± 30	EVANGELIS... 79	OMEG	10,16 $\pi^- p \rightarrow \bar{p} p n$	

<b><math>\rho_2</math>(1940)</b> $I^G(J^{PC}) = 1^+(2^{- -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1940 ± 40	155 ± 40	14 ANISOVICH	02	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$

<sup>14</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

<b><math>\omega_3</math>(1945)</b> $I^G(J^{PC}) = 0^-(3^{- -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1945 ± 20	115 ± 22	15 ANISOVICH	02B	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>15</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

<b><math>a_2</math>(1950)</b> $I^G(J^{PC}) = 1^-(2^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1950 $^{+30}_{-70}$	180 $^{+30}_{-70}$	16 ANISOVICH	01F	SPEC	1.96–2.41 $\bar{p}p$

<sup>16</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, and ANISOVICH 01F.

<b><math>\omega</math>(1960)</b> $I^G(J^{PC}) = 0^-(1^{- -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1960 ± 25	195 ± 60	17 ANISOVICH	02B	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>17</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

<b><math>b_1</math>(1960)</b> $I^G(J^{PC}) = 1^+(1^{+ -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1960 ± 35	230 ± 50	18 ANISOVICH	02	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$

<sup>18</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

<b><math>h_1</math>(1965)</b> $I^G(J^{PC}) = 0^-(1^{+ -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1965 ± 45	345 ± 75	19 ANISOVICH	02B	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>19</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

<b><math>f_1</math>(1970)</b> $I^G(J^{PC}) = 0^+(1^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1971 ± 15	240 ± 45	ANISOVICH	00J	SPEC	

<b>X(1970)</b> $I^G(J^{PC}) = ?^?(?^{??})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1970 ± 10	40 ± 20	CHLIAPNIK... 80	HBC	32	$K^+p \rightarrow 2K_S^0 2\pi X$

<b>X(1975)</b> $I^G(J^{PC}) = ?^?(?^{??})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
1973 ± 15	80	30	CASO	70	HBC 11.2 $\pi^- p \rightarrow \rho 2\pi$

<b><math>\omega_2</math>(1975)</b> $I^G(J^{PC}) = 0^-(2^{- -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1975 ± 20	175 ± 25	20 ANISOVICH	02B	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>20</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

<b><math>a_2</math>(1990)</b> $I^G(J^{PC}) = 1^-(2^{+ +})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
2050 ± 10 ± 40	190 ± 22 ± 100	18k	21 SCHEGELSKY	06	RVUE $\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
2003 ± 10 ± 19	249 ± 23 ± 32		LU	05	B852 18 $\pi^- p \rightarrow \omega\pi^-\pi^0 p$

<sup>21</sup> From analysis of L3 data at 183–209 GeV.

$\Gamma(\gamma\gamma) \Gamma(\pi^+\pi^-\pi^0) / \Gamma(\text{total})$					
VALUE (keV)	EVTs	DOCUMENT ID	TECN	COMMENT	
0.11 ± 0.04 ± 0.05	18k	22 SCHEGELSKY	06	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$

<sup>22</sup> From analysis of L3 data at 183–209 GeV.

<b><math>\rho</math>(2000)</b> $I^G(J^{PC}) = 1^+(1^{- -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2000 ± 30	260 ± 45	23 BUGG	04C	RVUE	Compilation
~ 1988	~ 244	HASAN	94	RVUE	$\bar{p}p \rightarrow \pi\pi$

<sup>23</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

<b><math>f_2</math>(2000)</b> $I^G(J^{PC}) = 0^+(2^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2001 ± 10	312 ± 32	ANISOVICH	00J	SPEC	
~ 1996	~ 134	HASAN	94	RVUE	$\bar{p}p \rightarrow \pi\pi$

<b>X(2000)</b> $I^G(J^{PC}) = 1^-(?^{?+})$						
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
1964 ± 35	225 ± 50	24 ARMSTRONG	93D	E760	$\bar{p}p \rightarrow 3\pi^0 \rightarrow 6\gamma$	
~ 2100	~ 500	24 ANTIPOV	77	CIBS	– 25 $\pi^- p \rightarrow p\pi^- \rho_3$	
2214 ± 15	355 ± 21	25 BALTAY	77	HBC	0 15 $\pi^- p \rightarrow \Delta^{++} 3\pi$	
2080 ± 40	340 ± 80	KALELKAR	75	HBC	+	15 $\pi^+ p \rightarrow p\pi^+ \rho_3$

<sup>24</sup> Cannot determine spin to be 3.

<sup>25</sup> BALTAY 77 favors  $J^P = ,3^+$ .

<b>X(2000)</b> $I^G(J^{PC}) = ?^?(4^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
1998 ± 3 ± 5	<15	VLADIMIRSK...03	SPEC	$\pi^- p \rightarrow K_S^0 K_S^0 MM$	

<b><math>\pi_2</math>(2005)</b> $I^G(J^{PC}) = 1^-(2^{- +})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
1974 ± 14 ± 83	341 ± 61 ± 139	145k	LU	05	B852 18 $\pi^- p \rightarrow \omega\pi^-\pi^0 p$
2005 ± 15	200 ± 40		ANISOVICH	01F	SPEC 2.0 $\bar{p}p \rightarrow 3\pi^0, \pi^0\eta, \pi^0\eta'$

<b><math>\eta</math>(2010)</b> $I^G(J^{PC}) = 0^+(0^{- +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2010 $^{+35}_{-60}$	270 ± 60	ANISOVICH	00J	SPEC	

<b><math>\pi_1</math>(2015)</b> $I^G(J^{PC}) = 1^-(1^{- +})$					
MASS (MeV)	WIDTH (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
2014 ± 20 ± 16	230 ± 32 ± 73	145k	LU	05	B852 18 $\pi^- p \rightarrow \omega\pi^-\pi^0 p$
2001 ± 30 ± 92	333 ± 52 ± 49	69k	KUHN	04	B852 18 $\pi^- p \rightarrow \eta\pi^+\pi^-\pi^- p$

<b><math>a_0</math>(2020)</b> $I^G(J^{PC}) = 1^-(0^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2025 ± 30	330 ± 75	ANISOVICH	99C	SPEC	

<b>X(2020)</b> $I^G(J^{PC}) = ?^?(?^{??})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2015 ± 3	10 ± 4	FERRER	99	RVUE	$\pi p \rightarrow p p\bar{p}\pi(\pi)$

<b><math>h_3</math>(2025)</b> $I^G(J^{PC}) = 0^-(3^{+ -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2025 ± 20	145 ± 30	26 ANISOVICH	02B	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>26</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

<b><math>b_3</math>(2030)</b> $I^G(J^{PC}) = 1^+(3^{+ -})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2032 ± 12	117 ± 11	27 ANISOVICH	02	SPEC	0.6–1.9 $p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$

<sup>27</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

<b><math>a_2</math>(2030)</b> $I^G(J^{PC}) = 1^-(2^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2030 ± 20	205 ± 30	28 ANISOVICH	01F	SPEC	1.96–2.41 $\bar{p}p$

<sup>28</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, and ANISOVICH 01F.

<b><math>a_3</math>(2030)</b> $I^G(J^{PC}) = 1^-(3^{+ +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2031 ± 12	150 ± 18	29 ANISOVICH	01F	SPEC	1.96–2.41 $\bar{p}p$

<sup>29</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, and ANISOVICH 01F.

<b><math>\eta_2</math>(2030)</b> $I^G(J^{PC}) = 0^+(2^{- +})$					
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT	
2030 ± 5 ± 15	205 ± 10 ± 15	ANISOVICH	00E	SPEC	

# Meson Particle Listings

## Further States

<b>B(<i>a</i><sub>2</sub>π)<sub>L=0</sub>/B(<i>a</i><sub>2</sub>π)<sub>L=2</sub></b>				
VALUE	DOCUMENT ID	TECN	COMMENT	
0.05±0.03	30 ANISOVICH	11	SPEC	0.9–1.94 <i>p</i> $\overline{p}$
30 Reanalysis of ADOMEIT 96 and ANISOVICH 00E.				

<b>B(<i>a</i><sub>0</sub>π)/B(<i>a</i><sub>2</sub>π)<sub>L=2</sub></b>				
VALUE	DOCUMENT ID	TECN	COMMENT	
0.10±0.08	31 ANISOVICH	11	SPEC	0.9–1.94 <i>p</i> $\overline{p}$
31 Reanalysis of ADOMEIT 96 and ANISOVICH 00E.				

<b>B(<i>f</i><sub>2</sub>η)/B(<i>a</i><sub>2</sub>π)<sub>L=2</sub></b>				
VALUE	DOCUMENT ID	TECN	COMMENT	
0.13±0.06	32 ANISOVICH	11	SPEC	0.9–1.94 <i>p</i> $\overline{p}$
32 Reanalysis of ADOMEIT 96 and ANISOVICH 00E.				

<b><i>f</i><sub>3</sub>(2050) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>+</sup>(3<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2048±8	213±34	ANISOVICH	00J	SPEC 2.0 <i>p</i> $\overline{p}$ → ηπ <sup>0</sup> π <sup>0</sup>

<b><i>f</i><sub>0</sub>(2060) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>+</sup>(0<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
~ 2050	~ 120	33 OAKDEN	94	RVUE 0.36–1.55 <i>p</i> $\overline{p}$ → ππ
~ 2060	~ 50	33 OAKDEN	94	RVUE 0.36–1.55 <i>p</i> $\overline{p}$ → ππ
33 See SEMENOV 99 and KLOET 96.				

<b>π(2070) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>−</sup>(0<sup>−+</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2070±35	310 <sup>+100</sup> <sub>−50</sub>	ANISOVICH	01F	SPEC 2.0 <i>p</i> $\overline{p}$ → 3π <sup>0</sup> , π <sup>0</sup> η, π <sup>0</sup> η′

<b>X(2075) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(?<sup>??</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2075±12±5	90±35±9	34 ABLIKIM	04J	BES2 <i>J</i> /ψ → <i>K</i> <sup>−</sup> <i>p</i> $\overline{\Lambda}$
34 From a fit in the region <i>M</i> <sub><i>p</i><math>\overline{\Lambda}</math></sub> − <i>M</i> <sub><i>p</i></sub> − <i>M</i> <sub>Λ</sub> < 150 MeV. <i>S</i> -wave in the <i>p</i> $\overline{\Lambda}$ system preferred. A similar near-threshold enhancement in the <i>p</i> $\overline{\Lambda}$ system is observed in <i>B</i> <sup>+</sup> → <i>p</i> $\overline{\Lambda}$ $\overline{D}^0$ by CHEN 11F.				

<b>X(2080) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(?<sup>??</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2080±10	110±20	KREYMER	80	STRC 13 π <sup>−</sup> <i>d</i> → <i>p</i> $\overline{p}$ <i>n</i> ( <i>n</i> <sub>5</sub> )

<b>X(2080) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(3<sup>−2</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2080±10	190±15	ROZANSKA	80	SPRK 18 π <sup>−</sup> <i>p</i> → <i>p</i> $\overline{p}$ <i>n</i>

<b><i>a</i><sub>1</sub>(2095) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>−</sup>(1<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	EVTS	DOCUMENT ID	TECN COMMENT
2096±17±121	451±41±81	69k	KUHN	04 B852 18 π <sup>−</sup> <i>p</i> → ηπ <sup>+</sup> π <sup>−</sup> π <sup>−</sup> <i>p</i>

<b>B(<i>a</i><sub>1</sub>(2095) → <i>f</i><sub>1</sub>(1285)π) / B(<i>a</i><sub>1</sub>(2095) → <i>a</i><sub>1</sub>(1260))</b>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
3.18±0.64	69k	KUHN	04	B852 18 π <sup>−</sup> <i>p</i> → ηπ <sup>+</sup> π <sup>−</sup> π <sup>−</sup> <i>p</i>

<b>η(2100) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>+</sup>(0<sup>−+</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	EVTS	DOCUMENT ID	TECN COMMENT
2103±50	187±75	586	35 BISELLO	89B DM2 <i>J</i> /ψ → 4πγ
35 ASTON 81B sees no peak, has 850 events in Ajinenko+Barth bins. ARESTOV 80 sees no peak.				

<b>X(2100) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(0<sup>??</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2100±40	250±40	ALDE	86D	GAM4 100 π <sup>−</sup> <i>p</i> → 2ηX

<b>X(2110) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>+</sup>(3<sup>−2</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2110±10	330±20	EVANGELIS...	79	OMEG 10,16 π <sup>−</sup> <i>p</i> → $\overline{p}$ <i>p</i> <i>n</i>

<b><i>f</i><sub>2</sub>(2140) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>+</sup>(2<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	EVTS	DOCUMENT ID	TECN COMMENT
2141±12	49±28	389	GREEN	86 MPSF 400 <i>pA</i> → 4 <i>K</i> X

<b>X(2150) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(2<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2150±10	260±10	ROZANSKA	80	SPRK 18 π <sup>−</sup> <i>p</i> → <i>p</i> $\overline{p}$ <i>n</i>

<b><i>a</i><sub>2</sub>(2175) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>−</sup>(2<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2175±40	310 <sup>+90</sup> <sub>−45</sub>	ANISOVICH	01F	SPEC 2.0 $\overline{p}$ <i>p</i> → 3π <sup>0</sup> , π <sup>0</sup> η, π <sup>0</sup> η′

<b>η(2190) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>+</sup>(0<sup>−+</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2190±50	850±100	BUGG	99	BES

<b>ω<sub>2</sub>(2195) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>−</sup>(2<sup>−−</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2195±30	225±40	36 ANISOVICH	02B	SPEC 0.6–1.9 <i>p</i> $\overline{p}$ → ωη, ωπ <sup>0</sup> π <sup>0</sup>
36 From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.				

<b>ω(2205) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>−</sup>(1<sup>−−</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2205±30	350±90	37 ANISOVICH	02B	SPEC 0.6–1.9 <i>p</i> $\overline{p}$ → ωη, ωπ <sup>0</sup> π <sup>0</sup>
37 From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.				

<b>X(2210) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(?<sup>??</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2210 <sup>+79</sup> <sub>−21</sub>	203 <sup>+437</sup> <sub>−87</sub>	EVANGELIS...	79B	OMEG 10 π <sup>−</sup> <i>p</i> → <i>K</i> <sup>+</sup> <i>K</i> <sup>−</sup> <i>n</i>

<b>X(2210) <i>I</i><sup>G</sup>(<i>JPC</i>) = ?<sup>?</sup>(?<sup>??</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2207±22	130	CASO	70	HBC 11.2 π <sup>−</sup> <i>p</i>

<b><i>h</i><sub>1</sub>(2215) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>−</sup>(1<sup>−+</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2215±40	325±55	38 ANISOVICH	02B	SPEC 0.6–1.9 <i>p</i> $\overline{p}$ → ωη, ωπ <sup>0</sup> π <sup>0</sup>
38 From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.				

<b><i>ρ</i><sub>2</sub>(2225) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>+</sup>(2<sup>−−</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2225±35	335 <sup>+100</sup> <sub>−50</sub>	39 ANISOVICH	02	SPEC 0.6–1.9 <i>p</i> $\overline{p}$ → ωπ <sup>0</sup> , ωηπ <sup>0</sup> , π <sup>+</sup> π <sup>−</sup>
39 From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.				

<b><i>ρ</i><sub>4</sub>(2230) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>+</sup>(4<sup>−−</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2230±25	210±30	40 ANISOVICH	02	SPEC 0.6–1.9 <i>p</i> $\overline{p}$ → ωπ <sup>0</sup> , ωηπ <sup>0</sup> , π <sup>+</sup> π <sup>−</sup>
40 From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.				

<b><i>b</i><sub>1</sub>(2240) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>+</sup>(1<sup>−+</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2240±35	320±85	41 ANISOVICH	02	SPEC 0.6–1.9 <i>p</i> $\overline{p}$ → ωπ <sup>0</sup> , ωηπ <sup>0</sup> , π <sup>+</sup> π <sup>−</sup>
41 From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.				

<b><i>f</i><sub>2</sub>(2240) <i>I</i><sup>G</sup>(<i>JPC</i>) = 0<sup>+</sup>(2<sup>++</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2240±15	241±30	42 ANISOVICH	00J	SPEC 1.92–2.41 <i>p</i> $\overline{p}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 2226	~ 226	HASAN	94	RVUE <i>p</i> $\overline{p}$ → ππ
42 From the combined analysis of ANISOVICH 99C, ANISOVICH 99F, ANISOVICH 99J, ANISOVICH 99K, and ANISOVICH 00B. See also ANISOVICH 12.				

<b><i>b</i><sub>3</sub>(2245) <i>I</i><sup>G</sup>(<i>JPC</i>) = 1<sup>+</sup>(3<sup>−+</sup>)</b>				
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2245±50	320±70	43 BUGG	04C	RVUE



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## Further States

<sup>43</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

$\eta_2(2250)$	$I^G(J^{PC}) = 0^+(2^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2248 ± 20	280 ± 20	ANISOVICH	00J	SPEC
2267 ± 14	290 ± 50	ANISOVICH	00J	SPEC

$\pi_4(2250)$	$I^G(J^{PC}) = 1^-(4^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2250 ± 15	215 ± 25	ANISOVICH	01F	SPEC 2.0 $\bar{p}p \rightarrow 3\pi^0, \pi^0\eta, \pi^0\eta'$

$\omega_4(2250)$	$I^G(J^{PC}) = 0^-(4^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2250 ± 30	150 ± 50	<sup>44</sup> ANISOVICH	02B	SPEC 0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>44</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

$\omega_5(2250)$	$I^G(J^{PC}) = 0^-(5^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2250 ± 70	320 ± 95	<sup>45</sup> BUGG	04	RVUE

<sup>45</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

$\omega_3(2255)$	$I^G(J^{PC}) = 0^-(3^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2255 ± 15	175 ± 30	<sup>46</sup> ANISOVICH	02B	SPEC 0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>46</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

$a_4(2255)$	$I^G(J^{PC}) = 1^-(4^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2237 ± 5	291 ± 12	UMAN	06	E835 5.2 $\bar{p}p \rightarrow \eta\eta\pi^0$
2255 ± 40	330 $^{+110}_{-50}$	<sup>47</sup> ANISOVICH	01F	SPEC 1.96–2.41 $\bar{p}p$

<sup>47</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, and ANISOVICH 01F.

$a_2(2255)$	$I^G(J^{PC}) = 1^-(2^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2255 ± 20	230 ± 15	<sup>48</sup> ANISOVICH	01G	SPEC 1.96–2.41 $\bar{p}p$

<sup>48</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, ANISOVICH 01F, and ANISOVICH 01G.

$X(2260)$	$I^G(J^{PC}) = 0^+(4^+?)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2260 ± 20	400 ± 100	EVANGELIS...	79	OMEG 10,16 $\pi^-p \rightarrow \bar{p}pn$

$\rho(2270)$	$I^G(J^{PC}) = 1^+(1^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2265 ± 40	325 ± 80	<sup>49</sup> ANISOVICH	02	SPEC 0.6–1.9 $p\bar{p} \rightarrow \omega\pi^0, \omega\eta\pi^0, \pi^+\pi^-$
2280 ± 50	440 ± 110	ATKINSON	85	OMEG 20–70 $\gamma p \rightarrow p\omega\pi^+\pi^-\pi^0$

<sup>49</sup> From the combined analysis of ANISOVICH 00J, ANISOVICH 01D, ANISOVICH 01E, and ANISOVICH 02.

$a_1(2270)$	$I^G(J^{PC}) = 1^-(1^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2270 $^{+55}_{-40}$	305 $^{+70}_{-40}$	ANISOVICH	01F	SPEC 2.0 $\bar{p}p \rightarrow 3\pi^0, \pi^0\eta, \pi^0\eta'$

$h_3(2275)$	$I^G(J^{PC}) = 0^-(3^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2275 ± 25	190 ± 45	<sup>50</sup> ANISOVICH	02B	SPEC 0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>50</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

$a_3(2275)$	$I^G(J^{PC}) = 1^-(3^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2275 ± 35	350 $^{+100}_{-50}$	<sup>51</sup> ANISOVICH	01G	SPEC 1.96–2.41 $\bar{p}p$

<sup>51</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99E, ANISOVICH 01F, and ANISOVICH 01G.

$\pi_2(2285)$	$I^G(J^{PC}) = 1^-(2^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2285 ± 20 ± 25	250 ± 20 ± 25	<sup>52</sup> ANISOVICH	11	SPEC 0.9–1.94 $p\bar{p}$

<sup>52</sup> Reanalysis of ADOMEIT 96 and ANISOVICH 00E.

$\omega_3(2285)$	$I^G(J^{PC}) = 0^-(3^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2278 ± 28	224 ± 50	<sup>53</sup> BUGG	04A	RVUE
2285 ± 60	230 ± 40	<sup>54</sup> ANISOVICH	02B	SPEC 0.6–1.9 $p\bar{p} \rightarrow \omega\eta, \omega\pi^0\pi^0$

<sup>53</sup> Partial wave analysis of the data on  $p\bar{p} \rightarrow \bar{\Lambda}\Lambda$  from BARNES 00.

<sup>54</sup> From the combined analysis of ANISOVICH 00D, ANISOVICH 01C, and ANISOVICH 02B.

$\omega(2290)$	$I^G(J^{PC}) = 0^-(1^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2290 ± 20	275 ± 35	<sup>55</sup> BUGG	04A	RVUE

<sup>55</sup> Partial wave analysis of the data on  $p\bar{p} \rightarrow \bar{\Lambda}\Lambda$  from BARNES 00.

$f_2(2295)$	$I^G(J^{PC}) = 0^+(2^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2293 ± 13	216 ± 37	<sup>56</sup> ANISOVICH	00J	SPEC 1.92–2.41 $p\bar{p}$

<sup>56</sup> From the combined analysis of ANISOVICH 99C, ANISOVICH 99F, ANISOVICH 99J, ANISOVICH 99K, and ANISOVICH 00B. See also ANISOVICH 12.

$f_3(2300)$	$I^G(J^{PC}) = 0^+(3^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2334 ± 25	200 ± 20	<sup>57</sup> BUGG	04A	RVUE

<sup>57</sup> Partial wave analysis of the data on  $p\bar{p} \rightarrow \bar{\Lambda}\Lambda$  from BARNES 00.

$f_1(2310)$	$I^G(J^{PC}) = 0^+(1^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2310 ± 60	255 ± 70	ANISOVICH	00J	SPEC

$\eta(2320)$	$I^G(J^{PC}) = 0^+(0^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2320 ± 15	230 ± 35	<sup>58</sup> ANISOVICH	00M	SPEC

<sup>58</sup> From the combined analysis of  $\bar{p}p \rightarrow \eta\eta\eta$  from ANISOVICH 00M and  $\bar{p}p \rightarrow \eta\pi^0\pi^0$  from ANISOVICH 00J.

$\eta_4(2330)$	$I^G(J^{PC}) = 0^+(4^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2328 ± 38	240 ± 90	ANISOVICH	00J	SPEC 2.0 $p\bar{p} \rightarrow \eta\pi^0\pi^0$

$\omega(2330)$	$I^G(J^{PC}) = 0^-(1^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2330 ± 30	435 ± 75	ATKINSON	88	OMEG 25–50 $\gamma p \rightarrow \rho^\pm \rho^0 \pi^\mp$

$X(2340)$	$I^G(J^{PC}) = ?^?(???)$			
MASS (MeV)	WIDTH (MeV)	EVTS	DOCUMENT ID	TECN COMMENT
2340 ± 20	180 ± 60	126	<sup>59</sup> BALTAY	75 HBC 15 $\pi^+p \rightarrow p5\pi$

<sup>59</sup> Dominant decay into  $\rho^0\rho^0\pi^+$ . BALTAY 78 finds confirmation in  $2\pi^+\pi^-2\pi^0$  events which contain  $\rho^+\rho^0\pi^0$  and  $2\rho^+\pi^-$ .

$\pi(2360)$	$I^G(J^{PC}) = 1^-(0^-+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2360 ± 25	300 $^{+100}_{-50}$	ANISOVICH	01F	SPEC 2.0 $\bar{p}p \rightarrow 3\pi^0, \pi^0\eta, \pi^0\eta'$

$X(2360)$	$I^G(J^{PC}) = ?^?(4^+?)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2360 ± 10	430 ± 30	ROZANSKA	80	SPRK 18 $\pi^-p \rightarrow p\bar{p}n$

$X(2440)$	$I^G(J^{PC}) = ?^?(5^-?)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2440 ± 10	310 ± 20	ROZANSKA	80	SPRK 18 $\pi^-p \rightarrow p\bar{p}n$

$X(2540)$	$I^G(J^{PC}) = 0^+(0^+)$			
MASS (MeV)	WIDTH (MeV)	DOCUMENT ID	TECN	COMMENT
2539 ± 14 $^{+38}_{-14}$	274 $^{+77+126}_{-61-163}$	UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$

$\Gamma(\gamma\gamma) \times B(K\bar{K})$				
VALUE (eV)	DOCUMENT ID	TECN	COMMENT	
40 $^{+9+17}_{-7-40}$	UEHARA	13	BELL	$\gamma\gamma \rightarrow K_S^0 K_S^0$

# Meson Particle Listings

## Further States

$X(2632) \quad I^G(J^{PC}) = ?^?(???)$					
<u>MASS (MeV)</u>	<u>WIDTH (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
2635.2±3.3		<sup>60</sup> EVDOKIMOV 04	SELX	$X(2632) \rightarrow D_s^+ \eta$	
2631.6±2.1	< 17	<sup>61</sup> EVDOKIMOV 04	SELX	$X(2632) \rightarrow D_s^0 K^+$	
<sup>60</sup> From a mass difference to $D_s^+$ of 666.9 ± 3.3 MeV.					
<sup>61</sup> From a mass difference to $D_s^0$ of 767.0 ± 2.0 MeV.					
<b><math>B(X(2632) \rightarrow D^0 K^+)/B(X(2632) \rightarrow D_s^+ \eta)</math></b>					
<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>		
0.14±0.06		<sup>62</sup> EVDOKIMOV 04	SELX		
<sup>62</sup> Possible interpretation of this decay pattern is discussed by YASUI 07.					

$X(2680) \quad I^G(J^{PC}) = ?^?(???)$					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
2676±27	150	CASO 70	HBC	$11.2 \pi^- p \rightarrow \rho^- \pi^+ \pi^- p$	

$X(2710) \quad I^G(J^{PC}) = ?^?(6^+?)$					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
2710±20	170 ± 40	ROZANSKA 80	SPRK	$18 \pi^- p \rightarrow p \overline{p} n$	

$X(2750) \quad I^G(J^{PC}) = ?^?(7^-?)$					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
2747±32	195 ± 75	DENNEY 83	LASS	$10 \pi^+ p \rightarrow K^+ K^- \pi^+ p$	

$f_0(3100) \quad I^G(J^{PC}) = 0^+(6^+ +)$					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
3100±100	700 ± 130	BINON 05	GAMS	$33 \pi^- p \rightarrow \eta \eta n$	

$X(3250) \quad I^G(J^{PC}) = ?^?(???)$ 3-Body Decays					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
3250±8±20	45 ± 18	ALEEV 93	BIS2	$X(3250) \rightarrow \Lambda \overline{p} K^+$	
3265±7±20	40 ± 18	ALEEV 93	BIS2	$X(3250) \rightarrow \overline{\Lambda} p K^-$	

$X(3250) \quad I^G(J^{PC}) = ?^?(???)$ 4-Body Decays					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
3245±8±20	25 ± 11	ALEEV 93	BIS2	$X(3250) \rightarrow \Lambda \overline{p} K^+ \pi^\pm$	
3250±9±20	50 ± 20	ALEEV 93	BIS2	$X(3250) \rightarrow \overline{\Lambda} p K^- \pi^\mp$	
3270±8±20	25 ± 11	ALEEV 93	BIS2	$X(3250) \rightarrow K_S^0 \rho \overline{p} K^\pm$	

$X(3350) \quad I^G(J^{PC}) = ?^?(???)$					
<i>MASS</i> (MeV)	<i>WIDTH</i> (MeV)	<i>EVTs</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
$3350^{+10}_{-20} \pm 20$	$70^{+40}_{-30} \pm 40$	50 ± 10	<sup>63</sup> GABYSHEV 06A	BELL	$B^- \rightarrow \Lambda_c^+ \overline{p} \pi^-$
<sup>63</sup> A similar enhancement in the $\Lambda_c^+ \overline{p}$ final state is also reported by BABAR collaboration in AUBERT 10H.					

### REFERENCES for Further States

ABLIKIM	13J	PR D87 032008	M. Ablikim <i>et al.</i>	(BES III Collab.)
UEHARA	13	PTEP 2013 123C01	S. Uehara <i>et al.</i>	(BELLE Collab.)
ANISOVICH	12	PR D85 014001	A.V. Anisovich <i>et al.</i>	
ANASHIN	11	PL B703 543	V.V. Anashin <i>et al.</i>	(KEDR Collab.)
ANISOVICH	11	EPJ C71 1511	A.V. Anisovich <i>et al.</i>	(LOQM, RAL, PNPI)
CHEN	11F	PR D84 071501	P. Chen <i>et al.</i>	(BELLE Collab.)
AUBERT	10H	PR D82 031102	B. AUBERT <i>et al.</i>	(BABAR Collab.)

ABRAAMYAN	09	PR C80 034001	Kh.U. Abraamyan <i>et al.</i>	
LIU	09	PR D79 071102	C. Liu <i>et al.</i>	(BELLE Collab.)
VLADIMIRSK...	08	PAN 71 2129	V.V. Vladimirsky <i>et al.</i>	(ITEP)
Translated from YAF 71 2166.				
VLADIMIRSK...	07	PAN 70 1706	V. Vladimirsky <i>et al.</i>	
Translated from YAF 70 1751.				
YASUI	07	PR D76 034009	S. Yasui, M. Oka	
ABLIKIM	06J	PRL 96 162002	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	06S	PRL 97 142002	M. Ablikim <i>et al.</i>	(BES Collab.)
GABYSHEV	06A	PRL 97 242001	N. Gabyshev <i>et al.</i>	(BELLE Collab.)
SCHEGELSKY	06A	EPJ A27 199	V.A. Schegelsky <i>et al.</i>	
SCHEGELSKY	06A	EPJ A27 207	V.A. Schegelsky <i>et al.</i>	
UMAN	06	PR D73 052009	I. Uman <i>et al.</i>	(FNAL E835)
VLADIMIRSK...	06	PAN 69 493	V.V. Vladimirsky <i>et al.</i>	(ITEP, Moscow)
Translated from YAF 69 515.				
BINON	05	PAN 68 960	F. Binon <i>et al.</i>	
Translated from YAF 68 998.				
GRIGOR'EV	05	PAN 68 1271	V.K. Grigor'ev <i>et al.</i>	(ITEP)
Translated from YAF 68 1324.				
LU	05	PRL 94 032002	M. Lu <i>et al.</i>	(BNL E852 Collab.)
ABLIKIM	04J	PRL 93 112002	M. Ablikim <i>et al.</i>	(BES Collab.)
BUGG	04	PL B595 556 (errat.)	D.V. Bugg	
BUGG	04A	EPJ C36 161	D.V. Bugg	
BUGG	04C	PRPL 397 257	D.V. Bugg	
EVDOKIMOV	04	PRL 93 242001	A.V. Evdokimov <i>et al.</i>	(SELEX Collab.)
KUHN	04	PL B595 109	J. Kuhn <i>et al.</i>	(BNL E852 Collab.)
ANISOVICH	03	EPJ A16 229	V.V. Anisovich <i>et al.</i>	
VLADIMIRSK...	03	PAN 66 700	V.V. Vladimirsky <i>et al.</i>	
Translated from YAF 66 729.				
ANISOVICH	02	PL B542 8	A.V. Anisovich <i>et al.</i>	
ANISOVICH	02B	PL B542 19	A.V. Anisovich <i>et al.</i>	
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
LINK	02K	PL B545 50	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ANISOVICH	01C	PL B507 23	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01D	PL B508 6	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01E	PL B513 281	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01F	PL B517 261	A.V. Anisovich <i>et al.</i>	
ANISOVICH	01G	PL B517 273	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00B	NP A662 319	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00D	PL B476 15	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00E	PL B477 19	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00I	PL B491 40	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00J	PL B491 47	A.V. Anisovich <i>et al.</i>	
ANISOVICH	00M	PL B496 145	A.V. Anisovich <i>et al.</i>	
BARNES	00	PR C62 055203	P.D. Barnes <i>et al.</i>	
FILIPPI	00	PL B495 284	A. Filippi <i>et al.</i>	(OBELIX Experiment)
VLADIMIRSKII	00	JETPL 72 486	V.V. Vladimirskii <i>et al.</i>	
Translated from ZETFP 72 698.				
ANISOVICH	99C	PL B452 173	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99E	PL B452 187	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99F	NP A651 253	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99J	PL B471 271	A.V. Anisovich <i>et al.</i>	
ANISOVICH	99K	PL B468 309	A.V. Anisovich <i>et al.</i>	
BUGG	99	PL B458 511	D.V. Bugg <i>et al.</i>	
FERRER	99	EPJ C10 249	A. Ferrer <i>et al.</i>	
SEMENOV	99	SPU 42 847	S.V. Semenov	
Translated from UFN 42 937.				
ADOMEIT	96	ZPHY C71 227	J. Adomeit <i>et al.</i>	(Crystal Barrel Collab.)
KLOET	96	PR D53 6120	W.M. Kloet, F. Myhrer	(RUTG, NORD)
PROKOSHKIN	96	SPD 41 247	V.D. Prokoshkin, V.D. Samoilenko	(SERP)
Translated from DANS 348 481.				
HASAN	94	PL B334 215	A. Hasan, D.V. Bugg	(LOQM)
OAKDEN	94	NP A574 731	M.N. Oakden, M.R. Pennington	(DURH)
ALEEV	93	PAN 56 1358	A.N. Aleev <i>et al.</i>	(BIS-2 Collab.)
Translated from YAF 56 100.				
ARMSTRONG	93D	PL B307 399	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
ALBRECHT	91F	ZPHY C50 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
CONDO	91	PR D43 2787	G.T. Condo <i>et al.</i>	(SLAC Hybrid Collab.)
BISELLO	89B	PR D39 701	G. Busetto <i>et al.</i>	(DM2 Collab.)
ATKINSON	88	ZPHY C38 535	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
DAFTARI	87	PRL 58 859	I.K. Daftari <i>et al.</i>	(SYRA)
ALDE	86D	NP B269 485	D.M. Alde <i>et al.</i>	(BELG, LAPP, SERP, CERN+)
BRIDGES	86D	PL B180 313	D.L. Bridges <i>et al.</i>	(SYRA, BNL, CASE+)
GREEN	86	PRL 56 1639	D.R. Green <i>et al.</i>	(FNAL, ARIZ, FSU+)
ATKINSON	85	ZPHY C29 333	M. Atkinson <i>et al.</i>	(BONN, CERN, GLAS+)
DENNEY	83	PR D28 2726	D.L. Denney <i>et al.</i>	(IOWA, MICH)
ASTON	81B	NP B189 205	D. Aston <i>et al.</i>	(BONN, CERN, EPOL, GLAS+)
ARESTOV	80	IHEP 80-165	Y.I. Arestov <i>et al.</i>	(SERP)
CHIAPNIK...	80	ZPHY C3 285	P.V. Chiapnikov <i>et al.</i>	(SERP, BRUX, MONS)
KREYMER	80	PR D22 36	A.E. Kreymer <i>et al.</i>	(IND, PURD, SLAC+)
ROZANSKA	80	NP B162 505	M. Rozanska <i>et al.</i>	(MPIM, CERN)
EVANGELIS...	79	NP B153 253	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
EVANGELIS...	79B	NP B154 381	C. Evangelista <i>et al.</i>	(BARI, BONN, CERN+)
BALTAY	78	PR D17 52	C. Baltay <i>et al.</i>	(COLU, BING)
ANTIPOV	77	NP B119 45	Y.M. Antipov <i>et al.</i>	(SERP, GEVA)
BALTAY	77	PRL 39 591	C. Baltay, C.V. Cautis, M. Kalelkar	(COLU)
BALTAY	75	PRL 35 891	C. Baltay <i>et al.</i>	(COLU, BING)
KALELKAR	75	Thesis Nevis 207	M.S. Kalelkar	(COLU)
CASO	70	LCN 3 707	C. Caso <i>et al.</i>	(GENO, HAMB, MILA, SACL)

**STRANGE MESONS**

$$(S = \pm 1, C = B = 0)$$

$$K^+ = u\bar{s}, K^0 = d\bar{s}, \bar{K}^0 = \bar{d}s, K^- = \bar{u}s, \text{ similarly for } K^{*'}s$$

 $K^\pm$ 

$$I(J^P) = \frac{1}{2}(0^-)$$

**THE CHARGED KAON MASS**

Revised 1994 by T.G. Trippe (LBNL).

The average of the six charged kaon mass measurements which we use in the Particle Listings is

$$m_{K^\pm} = 493.677 \pm 0.013 \text{ MeV } (S = 2.4), \quad (1)$$

where the error has been increased by the scale factor  $S$ . The large scale factor indicates a serious disagreement between different input data. The average before scaling the error is

$$m_{K^\pm} = 493.677 \pm 0.005 \text{ MeV},$$

$$\chi^2 = 22.9 \text{ for 5 D.F., Prob.} = 0.04\%, \quad (2)$$

where the high  $\chi^2$  and correspondingly low  $\chi^2$  probability further quantify the disagreement.

The main disagreement is between the two most recent and precise results,

$$m_{K^\pm} = 493.696 \pm 0.007 \text{ MeV} \quad \text{DENISOV 91}$$

$$m_{K^\pm} = 493.636 \pm 0.011 \text{ MeV } (S = 1.5) \quad \text{GALL 88}$$

$$\text{Average} = 493.679 \pm 0.006 \text{ MeV}$$

$$\chi^2 = 21.2 \text{ for 1 D.F., Prob.} = 0.0004\%, \quad (3)$$

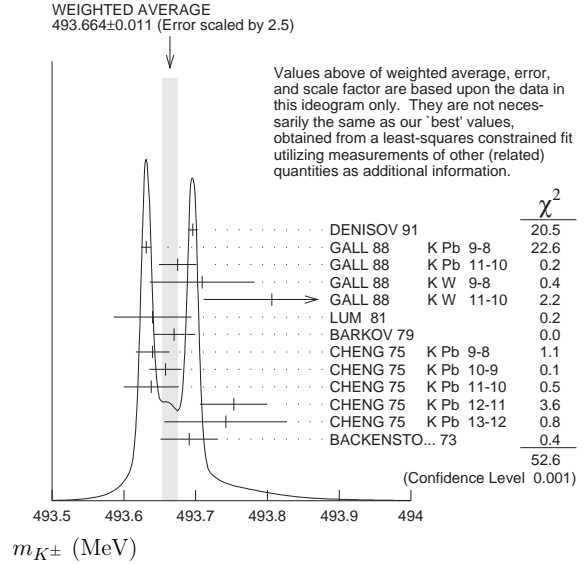
both of which are measurements of x-ray energies from kaonic atoms. Comparing the average in Eq. (3) with the overall average in Eq. (2), it is clear that DENISOV 91 and GALL 88 dominate the overall average, and that their disagreement is responsible for most of the high  $\chi^2$ .

The GALL 88 measurement was made using four different kaonic atom transitions,  $K^- \text{ Pb } (9 \rightarrow 8)$ ,  $K^- \text{ Pb } (11 \rightarrow 10)$ ,  $K^- \text{ W } (9 \rightarrow 8)$ , and  $K^- \text{ W } (11 \rightarrow 10)$ . The  $m_{K^\pm}$  values they obtain from each of these transitions is shown in the Particle Listings and in Fig. 1. Their  $K^- \text{ Pb } (9 \rightarrow 8)$   $m_{K^\pm}$  is below and somewhat inconsistent with their other three transitions. The average of their four measurements is

$$m_{K^\pm} = 493.636 \pm 0.007,$$

$$\chi^2 = 7.0 \text{ for 3 D.F., Prob.} = 7.2\%. \quad (4)$$

This is a low but acceptable  $\chi^2$  probability so, to be conservative, GALL 88 scaled up the error on their average by  $S=1.5$  to obtain their published error  $\pm 0.011$  shown in Eq. (3) above and used in the Particle Listings average.



**Figure 1:** Ideogram of  $m_{K^\pm}$  mass measurements. GALL 88 and CHENG 75 measurements are shown separately for each transition they measured.

The ideogram in Fig. 1 shows that the DENISOV 91 measurement and the GALL 88  $K^- \text{ Pb } (9 \rightarrow 8)$  measurement yield two well-separated peaks. One might suspect the GALL 88  $K^- \text{ Pb } (9 \rightarrow 8)$  measurement since it is responsible both for the internal inconsistency in the GALL 88 measurements and the disagreement with DENISOV 91.

**Table 1:**  $m_{K^\pm}$  averages for some combinations of Fig. 1 data.

$m_{K^\pm}$ (MeV)	$\chi^2$	D.F.	Prob. (%)	Measurements used
$493.664 \pm 0.004$	52.6	12	0.00005	all 13 measurements
$493.690 \pm 0.006$	10.1	10	43	no $K^- \text{ Pb } (9 \rightarrow 8)$
$493.687 \pm 0.006$	14.6	11	20	no GALL 88 $K^- \text{ Pb } (9 \rightarrow 8)$
$493.642 \pm 0.006$	17.8	11	8.6	no DENISOV 91

To see if the disagreement could result from a systematic problem with the  $K^- \text{ Pb } (9 \rightarrow 8)$  transition, we have separated the CHENG 75 data, which also used  $K^- \text{ Pb}$ , into its separate transitions. Figure 1 shows that the CHENG 75 and GALL 88  $K^- \text{ Pb } (9 \rightarrow 8)$  values are consistent, suggesting the possibility of a common effect such as contaminant nuclear  $\gamma$  rays near the  $K^- \text{ Pb } (9 \rightarrow 8)$  transition energy, although the CHENG 75 errors are too large to make a strong conclusion. The average of all 13 measurements has a  $\chi^2$  of 52.6 as shown in Fig. 1 and the first line of Table 1, yielding an unacceptable  $\chi^2$  probability of 0.00005%. The second line of Table 1 excludes both the GALL 88 and CHENG 75 measurements of the  $K^- \text{ Pb } (9 \rightarrow 8)$  transition and yields a  $\chi^2$  probability of 43%. The third [fourth] line of Table 1 excludes only the GALL 88  $K^- \text{ Pb } (9 \rightarrow 8)$  [DENISOV 91] measurement and yields a

# Meson Particle Listings

$K^\pm$

$\chi^2$  probability of 20% [8.6%]. Table 1 shows that removing both measurements of the  $K^-$  Pb ( $9 \rightarrow 8$ ) transition produces the most consistent set of data, but that excluding only the GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) transition or DENISOV 91 also produces acceptable probabilities.

Yu.M. Ivanov, representing DENISOV 91, has estimated corrections needed for the older experiments because of improved  $^{192}\text{Ir}$  and  $^{198}\text{Au}$  calibration  $\gamma$ -ray energies. He estimates that CHENG 75 and BACKENSTOSS 73  $m_{K^\pm}$  values could be raised by about 15 keV and 22 keV, respectively. With these estimated corrections, Table 1 becomes Table 2. The last line of Table 2 shows that if such corrections are assumed, then GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) is inconsistent with the rest of the data even when DENISOV 91 is excluded. Yu.M. Ivanov warns that these are rough estimates. Accordingly, we do not use Table 2 to reject the GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) transition, but we note that a future reanalysis of the CHENG 75 data could be useful because it might provide supporting evidence for such a rejection.

**Table 2:**  $m_{K^\pm}$  averages for some combinations of Fig. 1 data after raising CHENG 75 and BACKENSTOSS 73 values by 0.015 and 0.022 MeV respectively.

$m_{K^\pm}$ (MeV)	$\chi^2$	D.F.	Prob. (%)	Measurements used
$493.666 \pm 0.004$	53.9	12	0.00003	all 13 measurements
$493.693 \pm 0.006$	9.0	10	53	no $K^-$ Pb( $9 \rightarrow 8$ )
$493.690 \pm 0.006$	11.5	11	40	no GALL 88 $K^-$ Pb( $9 \rightarrow 8$ )
$493.645 \pm 0.006$	23.0	11	1.8	no DENISOV 91

The GALL 88 measurement uses a Ge semiconductor spectrometer which has a resolution of about 1 keV, so they run the risk of some contaminant nuclear  $\gamma$  rays. Studies of  $\gamma$  rays following stopped  $\pi^-$  and  $\Sigma^-$  absorption in nuclei (unpublished) do not show any evidence for contaminants according to GALL 88 spokesperson, B.L. Roberts. The DENISOV 91 measurement uses a crystal diffraction spectrometer with a resolution of 6.3 eV for radiation at 22.1 keV to measure the 4f-3d transition in  $K^-$   $^{12}\text{C}$ . The high resolution and the light nucleus reduce the probability for overlap by contaminant  $\gamma$  rays, compared with the measurement of GALL 88. The DENISOV 91 measurement is supported by their high-precision measurement of the 4d-2p transition energy in  $\pi^-$   $^{12}\text{C}$ , which is good agreement with the calculated energy.

While we suspect that the GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) measurements could be the problem, we are unable to find clear grounds for rejecting it. Therefore, we retain their measurement in the average and accept the large scale factor until further information can be obtained from new measurements and/or from reanalysis of GALL 88 and CHENG 75 data.

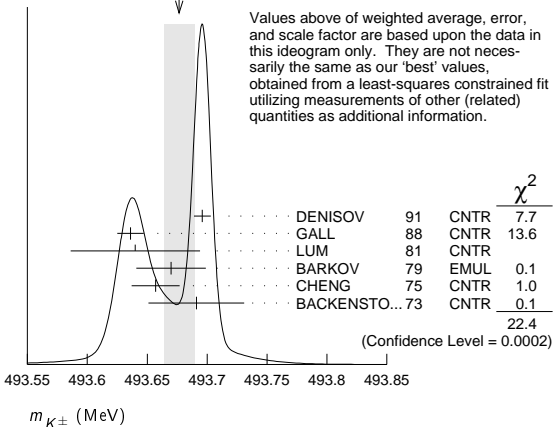
We thank B.L. Roberts (Boston Univ.) and Yu.M. Ivanov (Petersburg Nuclear Physics Inst.) for their extensive help in understanding this problem.

## $K^\pm$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>493.677 ± 0.016 OUR FIT</b>	Error includes scale factor of 2.8.				
<b>493.677 ± 0.013 OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.				
493.696 ± 0.007	<sup>1</sup> DENISOV	91	CNTR	—	Kaonic atoms
493.636 ± 0.011	<sup>2</sup> GALL	88	CNTR	—	Kaonic atoms
493.640 ± 0.054	LUM	81	CNTR	—	Kaonic atoms
493.670 ± 0.029	BARKOV	79	EMUL	±	e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup>
493.657 ± 0.020	<sup>2</sup> CHENG	75	CNTR	—	Kaonic atoms
493.691 ± 0.040	BACKENSTO..73	CNTR	—	—	Kaonic atoms
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
493.631 ± 0.007	GALL	88	CNTR	—	K <sup>-</sup> Pb (9 → 8)
493.675 ± 0.026	GALL	88	CNTR	—	K <sup>-</sup> Pb (11 → 10)
493.709 ± 0.073	GALL	88	CNTR	—	K <sup>-</sup> W (9 → 8)
493.806 ± 0.095	GALL	88	CNTR	—	K <sup>-</sup> W (11 → 10)
493.640 ± 0.022 ± 0.008	<sup>3</sup> CHENG	75	CNTR	—	K <sup>-</sup> Pb (9 → 8)
493.658 ± 0.019 ± 0.012	<sup>3</sup> CHENG	75	CNTR	—	K <sup>-</sup> Pb (10 → 9)
493.638 ± 0.035 ± 0.016	<sup>3</sup> CHENG	75	CNTR	—	K <sup>-</sup> Pb (11 → 10)
493.753 ± 0.042 ± 0.021	<sup>3</sup> CHENG	75	CNTR	—	K <sup>-</sup> Pb (12 → 11)
493.742 ± 0.081 ± 0.027	<sup>3</sup> CHENG	75	CNTR	—	K <sup>-</sup> Pb (13 → 12)

<sup>1</sup> Error increased from 0.0059 based on the error analysis in IVANOV 92.  
<sup>2</sup> This value is the authors' combination of all of the separate transitions listed for this paper.  
<sup>3</sup> The CHENG 75 values for separate transitions were calculated from their Table 7 transition energies. The first error includes a 20% systematic error in the noncircular contaminant shift. The second error is due to a  $\pm 5$  eV uncertainty in the theoretical transition energies.

WEIGHTED AVERAGE  
 $493.677 \pm 0.013$  (Error scaled by 2.4)



## $m_{K^+} - m_{K^-}$

Test of CPT.

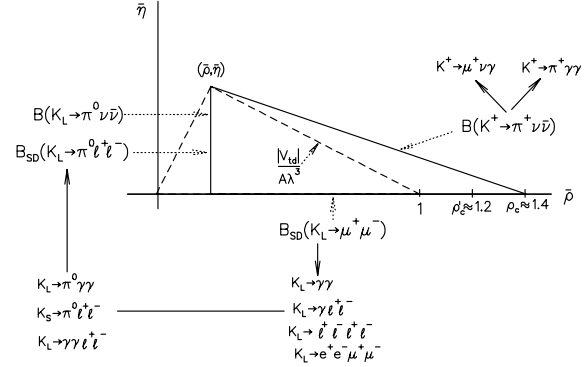
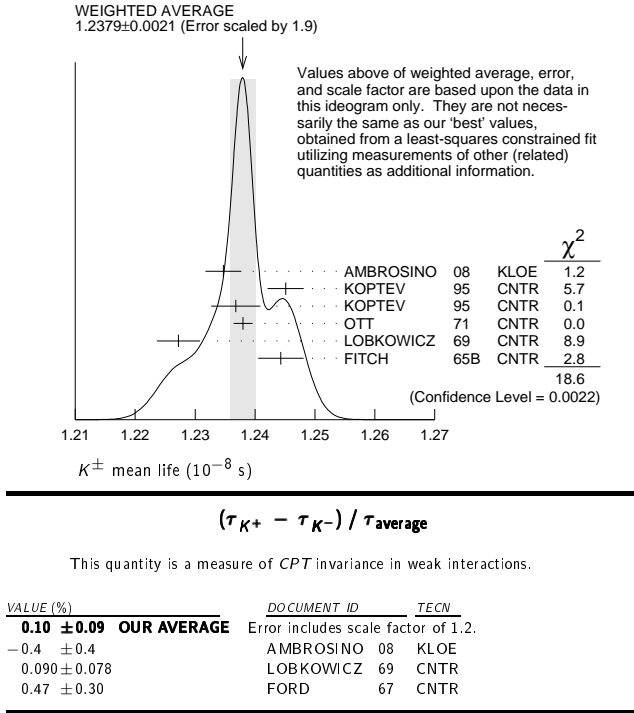
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG
<b><math>-0.032 \pm 0.090</math></b>	1.5 M	<sup>1</sup> FORD	72	ASP K ±

<sup>1</sup> FORD 72 uses  $m_{\pi^+} - m_{\pi^-} = +28 \pm 70$  keV.

## $K^\pm$ MEAN LIFE

VALUE ( $10^{-8}$ s)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1.2380 \pm 0.0020</math> OUR FIT</b>	Error includes scale factor of 1.8.				
<b><math>1.2379 \pm 0.0021</math> OUR AVERAGE</b>	Error includes scale factor of 1.9. See the ideogram below.				
$1.2347 \pm 0.0030$	15 M	<sup>1</sup> AMBROSINO	08	KLOE	± $\phi \rightarrow K^+ K^-$
$1.2451 \pm 0.0030$	250k	KOPTEV	95	CNTR	K at rest, U target
$1.2368 \pm 0.0041$	150k	KOPTEV	95	CNTR	K at rest, Cu target
$1.2380 \pm 0.0016$	3 M	OTT	71	CNTR	+ K at rest
$1.2272 \pm 0.0036$		LOBKOWICZ	69	CNTR	+ K in flight
$1.2443 \pm 0.0038$		FITCH	65B	CNTR	+ K at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.2415 \pm 0.0024$	400k	<sup>2</sup> KOPTEV	95	CNTR	K at rest
$1.221 \pm 0.011$		FORD	67	CNTR	±
$1.231 \pm 0.011$		BOYARSKI	62	CNTR	+

<sup>1</sup> Result obtained by averaging the decay length and decay time analyses taking correlations into account.  
<sup>2</sup> KOPTEV 95 report this weighted average of their U-target and Cu-target results, where they have weighted by  $1/\sigma$  rather than  $1/\sigma^2$ .



**Figure 1:** Role of rare kaon decays in determining the unitarity triangle. The solid arrows point to auxiliary modes needed to interpret the main results, or potential backgrounds to them.

## RARE KAON DECAYS

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**A. Introduction:** There are several useful reviews on rare kaon decays and related topics [1–17]. Activity in rare kaon decays can be divided roughly into four categories:

1. Searches for explicit violations of the Standard Model
2. Measurements of Standard Model parameters
3. Searches for direct *CP* violation
4. Studies of strong interactions at low energy.

The paradigm of Category 1 is the lepton flavor violating decay  $K_L \rightarrow \mu e$ . Category 2 includes processes such as  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ , which is sensitive to CKM parameters. Much of the interest in Category 3 is focused on the decays  $K_L \rightarrow \pi^0 \ell \ell$ , where  $\ell \equiv e, \mu, \nu$ . Category 4 includes reactions like  $K^+ \rightarrow \pi^+ \ell^+ \ell^-$  which constitute a testing ground for the ideas of chiral perturbation theory. Category 4 also includes  $K_L \rightarrow \pi^0 \gamma \gamma$  and  $K_L \rightarrow \ell^+ \ell^- \gamma$ . The former is important in understanding a *CP*-conserving contribution to  $K_L \rightarrow \pi^0 \ell^+ \ell^-$ , whereas the latter could shed light on long distance contributions to  $K_L \rightarrow \mu^+ \mu^-$ .

The interplay between Categories 2–4 can be illustrated in Fig. 1. The modes  $K \rightarrow \pi \nu \bar{\nu}$  are the cleanest ones theoretically. They can provide accurate determinations of certain CKM parameters (shown in the figure). In combination with alternate determinations of these parameters, they also constrain new interactions. The modes  $K_L \rightarrow \pi^0 e^+ e^-$ ,  $K_L \rightarrow \pi^0 \mu^+ \mu^-$  and  $K_L \rightarrow \mu^+ \mu^-$  are also sensitive to CKM parameters. However, they suffer from a number of hadronic uncertainties that can be addressed, at least in part, through a systematic study of the additional modes indicated in the figure.

**B. Explicit violations of the Standard Model:** Much activity has focussed on searches for lepton flavor violation (LFV). This is motivated by the fact that many extensions of the minimal Standard Model violate lepton flavor and by the potential to access very high energy scales. For example, the tree-level exchange of a LFV vector boson of mass  $M_X$  that couples to left-handed fermions with electroweak strength and without mixing angles yields  $B(K_L \rightarrow \mu e) = 4.7 \times 10^{-12} (148 \text{ TeV}/M_X)^4$  [4]. This simple dimensional analysis may be used to read from Table 1 that the reaction  $K_L \rightarrow \mu e$  is already probing scales of over 100 TeV. Table 1 summarizes the present experimental situation vis-à-vis LFV. The decays  $K_L \rightarrow \mu^\pm e^\mp$  and  $K^+ \rightarrow \pi^+ e^\mp \mu^\pm$  (or  $K_L \rightarrow \pi^0 e^\mp \mu^\pm$ ) provide complementary information on potential family number violating interactions, since the former is sensitive to parity-odd couplings and the latter is sensitive to parity-even couplings. Limits on certain lepton-number violating kaon decays also exist, some recent ones being those of Refs. [18,19,20]. Related searches in  $\mu$  and  $\tau$  processes are discussed in our section “Tests of Conservation Laws.”

**Table 1:** Searches for lepton flavor violation in *K* decay

Mode	90% CL upper limit	Exp’t	Yr./Ref.
$K^+ \rightarrow \pi^+ e^- \mu^+$	$1.2 \times 10^{-11}$	BNL-865	2005/Ref. 21
$K^+ \rightarrow \pi^+ e^+ \mu^-$	$5.2 \times 10^{-10}$	BNL-865	2000/Ref. 18
$K_L \rightarrow \mu e$	$4.7 \times 10^{-12}$	BNL-871	1998/Ref. 22
$K_L \rightarrow \pi^0 e \mu$	$7.6 \times 10^{-11}$	KTeV	2008/Ref. 23
$K_L \rightarrow \pi^0 \pi^0 e \mu$	$1.7 \times 10^{-10}$	KTeV	2008/Ref. 23

Physics beyond the SM is also pursued through the search for  $K^+ \rightarrow \pi^+ X^0$ , where  $X^0$  is a new light particle. The searches cover both long-lived particles (*e.g.*, hyperphoton, axion, familon, *etc.*), and short lived ones that decay to muon, electron or photon pairs. The 90% CL upper limit on  $K^+ \rightarrow \pi^+ X^0$  is  $7.3 \times 10^{-11}$  [24] for the case of massless  $X^0$ ; additional results as a function of the  $X^0$  mass can be found in [37].

# Meson Particle Listings

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Recently these limits have been reinterpreted in connection with a dark photon [25] or dark Z [26]. Such vectors have also been sought in their  $e^+e^-$  decay mode by NA48 [27]. Additional bounds for a short lived pseudoscalar  $X^0$  decaying to muons or photons are  $B(K_L \rightarrow \pi^0 \pi^0 \mu^+ \mu^-) < 1 \times 10^{-10}$  [28] and  $B(K_L \rightarrow \pi^0 \pi^0 \gamma \gamma) < 2.4 \times 10^{-7}$  [29].

## C. Measurements of Standard Model parameters:

In the SM, the decay  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  is dominated by one-loop diagrams with top-quark intermediate states and long-distance contributions are known to be quite small [2,30]. This permits a precise calculation of this rate in terms of SM parameters. Studies of this process are thus motivated by the possibility of detecting non-SM physics when comparing with the results of global fits [31,32].

BNL-787 observed two candidate events [33,34] in the clean high  $\pi^+$  momentum and one event [35] in the low-momentum region. The successor experiment BNL-949 observed one more in the high-momentum region [24] and three more in the low-momentum region [36] yielding a branching ratio of  $(1.73^{+1.15}_{-1.05}) \times 10^{-10}$  [37]. A subsequent experiment, NA62, with a sensitivity goal of  $\sim 10^{-12}$ /event was proposed [38] at CERN in 2005. It was approved and ran with a partial detector in autumn 2012, followed by a successful commissioning run in the fall of 2014. The first physics run started in the summer of 2015. The NA62 experiment will be the first one performed with kaon decays in flight. In the future, this mode may provide grounds for precision tests of flavor dynamics [40].

The branching ratio can be written in a compact form that exhibits the different ingredients that go into the calculation [41],

$$B(K^+ \rightarrow \pi^+ \nu \bar{\nu}(\gamma)) = \kappa_+(1 + \Delta_{\text{EM}}) \left[ \left( \frac{\text{Im}(V_{ts}^* V_{td})}{\lambda^5} X_t \right)^2 + \left( \frac{\text{Re}(V_{cs}^* V_{cd})}{\lambda} (P_c + \delta P_{c,u}) + \frac{\text{Re}(V_{ts}^* V_{td})}{\lambda^5} X_t \right)^2 \right]. \quad (1)$$

The parameters in Eq. (1) incorporate the *a priori* unknown hadronic matrix element in terms of the very well-measured  $K_{\ell 3}$  rate [2] in  $\kappa_+$ ; long distance QED corrections in  $\Delta_{\text{EM}}$  [43]; the Inami-Lim function for the short distance top-quark contribution [44] including NLO QCD corrections [45] and the two-loop electroweak correction [41], all in  $X_t$ ; and the charm-quark contributions due to short distance effects including NNLO QCD corrections [46] and NLO electroweak corrections via  $P_c$  [47], as well as certain long distance effects via  $\delta P_{c,u}$  [48]. An interesting approximate way to cast this result in terms of the CKM parameters  $\lambda$ ,  $V_{cb}$ ,  $\bar{\rho}$  and  $\bar{\eta}$  (see our Section on “The Cabibbo-Kobayashi-Maskawa mixing matrix”) [11] is:

$$B(K^+ \rightarrow \pi^+ \nu \bar{\nu}) \approx 1.6 \times 10^{-5} |V_{cb}|^4 [\sigma \bar{\eta}^2 + (\rho_c - \bar{\rho})^2], \quad (2)$$

where  $\rho_c \approx 1.45$  and  $\sigma \equiv 1/(1 - \frac{1}{2}\lambda^2)^2$ . Thus,  $B(K^+ \rightarrow \pi^+ \nu \bar{\nu})$  determines an ellipse in the  $\bar{\rho}$ ,  $\bar{\eta}$  plane with center  $(\rho_c, 0)$  and semiaxes  $\approx \frac{1}{|V_{cb}|^2} \sqrt{\frac{B(K^+ \rightarrow \pi^+ \nu \bar{\nu})}{1.6 \times 10^{-5}}}$  and  $\frac{1}{\sigma |V_{cb}|^2} \sqrt{\frac{B(K^+ \rightarrow \pi^+ \nu \bar{\nu})}{1.6 \times 10^{-5}}}$ . A recent numerical study leads to a predicted branching ratio

$(7.81^{+0.80}_{-0.71} \pm 0.29) \times 10^{-11}$  [41], near the lower end of the measurement of BNL-787 and 949. However, parametric uncertainty in the CKM angles can result in numbers that differ from this one by up to 10% [42].

Modes with an extra pion,  $K \rightarrow \pi \pi \nu \bar{\nu}$ , could also be used in the extraction of CKM parameters as they are also dominated by short distance contributions [49]. However, they occur at much lower rates with branching ratios of order  $10^{-13}$ , and the current best bound from KEK-391a is  $B(K_L \rightarrow \pi^0 \pi^0 \nu \bar{\nu}) < 8.1 \times 10^{-7}$  at 90% CL [50]. There is also an older bound of  $B(K^+ \rightarrow \pi^+ \pi^0 \nu \bar{\nu}) < 4.3 \times 10^{-5}$  at 90% CL [51] from BNL-787.

The decay  $K_L \rightarrow \mu^+ \mu^-$  also has a short distance contribution sensitive to the CKM parameter  $\bar{\rho}$ , given by [11]:

$$B_{\text{SD}}(K_L \rightarrow \mu^+ \mu^-) \approx 2.7 \times 10^{-4} |V_{cb}|^4 (\rho'_c - \bar{\rho})^2 \quad (3)$$

where  $\rho'_c$  depends on the charm quark mass and is approximately 1.2. This decay, however, is dominated by a long-distance contribution from a two-photon intermediate state. The absorptive (imaginary) part of the long-distance component is determined by the measured rate for  $K_L \rightarrow \gamma \gamma$  to be  $B_{\text{abs}}(K_L \rightarrow \mu^+ \mu^-) = (6.64 \pm 0.07) \times 10^{-9}$ ; and it almost completely saturates the observed rate  $B(K_L \rightarrow \mu^+ \mu^-) = (6.84 \pm 0.11) \times 10^{-9}$  [52]. The difference between the observed rate and the absorptive component can be attributed to the (coherent) sum of the short-distance amplitude and the real part of the long-distance amplitude. The latter cannot be derived directly from experiment [53], but can be estimated with certain assumptions [54,55].

The decay  $K_L \rightarrow e^+ e^-$  is completely dominated by long distance physics and is easier to estimate. The result,  $B(K_L \rightarrow e^+ e^-) \sim 9 \times 10^{-12}$  [53,56], is in good agreement with the BNL-871 measurement,  $(8.7^{+5.7}_{-4.1}) \times 10^{-12}$  [57].

The mode  $K_S \rightarrow \mu^+ \mu^-$  also has a short distance contribution proportional to the square of the CKM parameter  $\bar{\eta}$  entering at the  $10^{-13}$  level [15] as well as long distance contributions which arise in this case from a two photon intermediate state and result in a rate  $B(K_S \rightarrow \mu^+ \mu^-)_{\text{LD}} = 5.1 \times 10^{-12}$  [15]. A 95% (90%) c.l. limit  $B(K_S \rightarrow \mu^+ \mu^-) < 11(9) \times 10^{-9}$  was obtained by LHCb [58].

**D. Searches for direct CP violation:** The mode  $K_L \rightarrow \pi^0 \nu \bar{\nu}$  is dominantly CP-violating and free of hadronic uncertainties [2,59,60]. In the Standard Model, this mode is dominated by an intermediate top-quark state and does not suffer from the small uncertainty associated with the charm-quark intermediate state that affects the mode  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ . The branching ratio is given by Ref. 11:

$$B(K_L \rightarrow \pi^0 \nu \bar{\nu}) = \kappa_L \left( \frac{\text{Im}(V_{ts}^* V_{td})}{\lambda^5} X_t \right)^2 \approx 7.6 \times 10^{-5} |V_{cb}|^4 \bar{\eta}^2. \quad (4)$$

The hadronic matrix element can be related to that measured in  $K_{\ell 3}$  decay and is parameterized in  $\kappa_L$ . A recent numerical evaluation leads to a predicted branching ratio  $(2.43^{+0.40}_{-0.37} \pm$

$0.06) \times 10^{-11}$  [41]. As noted for the charged kaon mode, parametric uncertainty in the CKM angles can result in a central value that differs from this one by up to almost 20% [42]. The 90% CL bound on  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  provides a nearly model-independent bound  $B(K_L \rightarrow \pi^0 \nu \bar{\nu}) < 1.46 \times 10^{-9}$  [61]. KEK-391a, which took data in 2004 and 2005, has published a 90% CL upper bound of  $B(K_L \rightarrow \pi^0 \nu \bar{\nu}) \leq 2.6 \times 10^{-8}$  [62]. The KOTO experiment at J-PARC [63], whose initial goal is to observe it, had a short physics run in the spring of 2013, reaching a single event sensitivity of  $1.29 \times 10^{-8}$  [65], and resuming the data taking in May 2015. It was pointed out in a recent paper that the above Grossman-Nir bound [61] on the three body decay  $K_L \rightarrow \pi^0 \nu \bar{\nu}$  doesn't necessarily apply to two body decays such as  $K_L \rightarrow \pi^0 X^0$ , so that KOTO may be interesting for new physics searches at the current sensitivity level [64].

There has been much theoretical work on possible contributions to rare  $K$  decays beyond the SM. A comprehensive discussion of these can be found in Refs. [14] and [66].

The decay  $K_L \rightarrow \pi^0 e^+ e^-$  also has sensitivity to the CKM parameter  $\eta$  through its  $CP$ -violating component. There are both direct and indirect  $CP$ -violating amplitudes that can interfere. The direct  $CP$ -violating amplitude is short distance dominated and has been calculated in detail within the SM [8]. The indirect  $CP$ -violating amplitude can be inferred from a measurement of  $K_S \rightarrow \pi^0 e^+ e^-$ . The complete  $CP$ -violating contribution to the rate can be written as [67,68]:

$$B_{CPV} \approx 10^{-12} \left[ 15.7 |a_S|^2 \pm 1.4 \left( \frac{|V_{cb}|^2 \eta}{10^{-4}} \right) |a_S| + 0.12 \left( \frac{|V_{cb}|^2 \eta}{10^{-4}} \right)^2 \right] \quad (5)$$

where the three terms correspond to the indirect  $CP$  violation, the interference, and the direct  $CP$  violation respectively. The parameter  $a_S$  has been extracted by NA48 from a measurement of the decay  $K_S \rightarrow \pi^0 e^+ e^-$  with the result  $|a_S| = 1.06_{-0.21}^{+0.26} \pm 0.07$  [69], as well as from a measurement of the decay  $K_S \rightarrow \pi^0 \mu^+ \mu^-$  with the result  $|a_S| = 1.54_{-0.32}^{+0.40} \pm 0.06$  [70]. With current constraints on the CKM parameters, and assuming a positive sign for the interference term [68,71], this implies that  $B_{CPV}(K_L \rightarrow \pi^0 e^+ e^-) \approx (3.1 \pm 0.9) \times 10^{-11}$ , and that the indirect  $CP$  violation is larger than the direct  $CP$  violation. The complete  $CP$  violating amplitude for the related mode  $K_L \rightarrow \pi^0 \mu^+ \mu^-$  is predicted to be  $B_{CPV}(K_L \rightarrow \pi^0 \mu^+ \mu^-) \approx (1.4 \pm 0.5) \times 10^{-11}$  [72,15].

$K_L \rightarrow \pi^0 e^+ e^-$  also has a  $CP$ -conserving component dominated by a two-photon intermediate state. This component can be decomposed into an absorptive and a dispersive part. The absorptive part can be extracted from the measurement of the low  $m_{\gamma\gamma}$  region of the  $K_L \rightarrow \pi^0 \gamma \gamma$  spectrum. The rate and the shape of the distribution  $d\Gamma/dm_{\gamma\gamma}$  in  $K_L \rightarrow \pi^0 \gamma \gamma$  are well described in chiral perturbation theory in terms of three (*a priori*) unknown parameters [73,74].

Both KTeV and NA48 have studied the mode  $K_L \rightarrow \pi^0 \gamma \gamma$ , reporting similar results. KTeV finds  $B(K_L \rightarrow \pi^0 \gamma \gamma) = (1.29 \pm 0.03_{\text{stat}} \pm 0.05_{\text{sys}}) \times 10^{-6}$  [75], while NA48 finds  $B(K_L \rightarrow \pi^0 \gamma \gamma) = (1.36 \pm 0.03_{\text{stat}} \pm 0.03_{\text{sys}} \pm 0.03_{\text{norm}}) \times 10^{-6}$  [76]. Both experiments are consistent with a negligible rate in the low  $m_{\gamma\gamma}$  region, suggesting a very small  $CP$ -conserving component  $B_{CP}(K_L \rightarrow \pi^0 e^+ e^-) \sim \mathcal{O}(10^{-13})$  [68,74,76]. There remains some model dependence in the estimate of the dispersive part of the  $CP$ -conserving  $K_L \rightarrow \pi^0 e^+ e^-$  [68].

The related process,  $K_L \rightarrow \pi^0 \gamma e^+ e^-$ , is potentially an additional background in some region of phase space [77]. This process has been observed with a branching ratio of  $(1.62 \pm 0.14_{\text{stat}} \pm 0.09_{\text{sys}}) \times 10^{-8}$  [78].

The decay  $K_L \rightarrow \gamma \gamma e^+ e^-$  constitutes the dominant background to  $K_L \rightarrow \pi^0 e^+ e^-$ . It was first observed by BNL-845 [79], and subsequently confirmed with a much larger sample by KTeV [80]. It has been estimated that this background will enter at about the  $10^{-10}$  level [81,82], comparable to or larger than the signal level. Because of this, the observation of  $K_L \rightarrow \pi^0 e^+ e^-$  at the SM level will depend on background subtraction with good statistics. Possible alternative strategies are discussed in Ref. 68 and references cited therein.

The 90% CL upper bound for the process  $K_L \rightarrow \pi^0 e^+ e^-$  is  $2.8 \times 10^{-10}$  [82]. For the closely related muonic process, the published upper bound is  $B(K_L \rightarrow \pi^0 \mu^+ \mu^-) \leq 3.8 \times 10^{-10}$  [83], compared with the SM prediction of  $(1.5 \pm 0.3) \times 10^{-11}$  [72] (assuming positive interference between the direct- and indirect- $CP$  violating components).

A study of  $K_L \rightarrow \pi^0 \mu^+ \mu^-$  has indicated that it might be possible to extract the direct  $CP$ -violating contribution by a joint study of the Dalitz plot variables and the components of the  $\mu^+$  polarization [84]. The latter tends to be quite substantial so that large statistics may not be necessary.

Combined information from the two  $K_L \rightarrow \pi^0 \ell^+ \ell^-$  modes complements the  $K \rightarrow \pi \nu \bar{\nu}$  measurements in constraining physics beyond the SM [85].

#### E. Other long distance dominated modes:

The decays  $K^+ \rightarrow \pi^+ \ell^+ \ell^-$  ( $\ell = e$  or  $\mu$ ) have received considerable attention. The rate and spectrum have been measured for both the electron and muon modes [86,87,20].

The measurements have been used to exclude new physics such as a dark photon [25]. Ref. 67 has proposed a parametrization inspired by chiral perturbation theory, which provides a successful description of data but indicates the presence of large corrections beyond leading order. More work is needed to fully understand the origin of these large corrections. NA62 has now also observed the mode  $K^+ \rightarrow \pi^+ \pi^0 e^+ e^-$  [88] studied in [89].

The decay  $K^+ \rightarrow \pi^+ \gamma \gamma$  can be predicted in terms of one unknown parameter to leading order in  $\chi$ PT resulting in a correlation between the rate and the diphoton mass spectrum [90]. Certain important corrections at the next order are also known [91]. The rate was first measured by E787 [92], more recently NA48/2 [93] has obtained a more precise result with 6% error along with the corresponding spectrum fits. The most recent, and precise, result is from NA62 based on a sample

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of 232 events [94] and is still insufficient to distinguish between the leading order and next order  $\chi$ PT parametrizations.

Much information has been recorded by KTeV and NA48 on the rates and spectrum for the Dalitz pair conversion modes  $K_L \rightarrow \ell^+ \ell^- \gamma$  [95,96], and  $K_L \rightarrow \ell^+ \ell^- \ell'^+ \ell'^-$  for  $\ell, \ell' = e$  or  $\mu$  [19,97–99]. All these results are used to test hadronic models and could further our understanding of the long distance component in  $K_L \rightarrow \mu^+ \mu^-$ .

## References

- D. Bryman, Int. J. Mod. Phys. **A4**, 79 (1989).
- J. Hagelin and L. Littenberg, Prog. in Part. Nucl. Phys. **23**, 1 (1989).
- L. Littenberg and G. Valencia, Ann. Rev. Nucl. and Part. Sci. **43**, 729 (1993).
- J. Ritchie and S. Wojcicki, Rev. Mod. Phys. **65**, 1149 (1993).
- B. Winstein and L. Wolfenstein, Rev. Mod. Phys. **65**, 1113 (1993).
- G. D'Ambrosio *et al.*, *Radiative Non-Leptonic Kaon Decays*, in The DAΦNE Physics Handbook (second edition), eds. L. Maiani, G. Pancheri, and N. Paver (Frascati), Vol. I, 265 (1995).
- A. Pich, Rept. on Prog. in Phys. **58**, 563 (1995).
- G. Buchalla, A.J. Buras, and M.E. Lautenbacher, Rev. Mod. Phys. **68**, 1125 (1996).
- G. D'Ambrosio and G. Isidori, Int. J. Mod. Phys. **A13**, 1 (1996).
- P. Buchholz and B. Renk Prog. in Part. Nucl. Phys. **39**, 253 (1997).
- A.J. Buras and R. Fleischer, TUM-HEP-275-97, [hep-ph/9704376](#), *Heavy Flavours II*, World Scientific, eds. A.J. Buras and M. Lindner (1997), 65–238.
- A.J. Buras, TUM-HEP-349-99, Lectures at Lake Louise Winter Institute: Electroweak Physics, Lake Louise, Alberta, Canada, 14–20 Feb. 1999.
- A.R. Barker and S.H. Kettell, Ann. Rev. Nucl. and Part. Sci. **50**, 249 (2000).
- A.J. Buras, F. Schwab, and S. Uhlig, Rev. Mod. Phys. **80**, 965 (2008).
- V. Cirigliano *et al.*, Rev. Mod. Phys. **84**, 399 (2012).
- D. Bryman *et al.*, Ann. Rev. Nucl. and Part. Sci. **61**, 331 (2011).
- T.K. Komatsubara, Prog. in Part. Nucl. Phys. **67**, 995 (2012).
- R. Appel *et al.*, Phys. Rev. Lett. **85**, 2877 (2000).
- A. Alavi-Harati *et al.*, Phys. Rev. Lett. **90**, 141801 (2003).
- J.R. Batley *et al.*, Phys. Lett. **B697**, 107 (2011).
- A. Sher *et al.*, Phys. Rev. **D72**, 012005 (2005).
- D. Ambrose *et al.*, Phys. Rev. Lett. **81**, 5734 (1998).
- E. Abouzaid *et al.*, Phys. Rev. Lett. **100**, 131803 (2008).
- V.V. Anisimovsky *et al.*, Phys. Rev. Lett. **93**, 031801 (2004).
- M. Pospelov, Phys. Rev. **D80**, 095002 (2009).
- H. Davoudiasl, H.S. Lee, and W.J. Marciano, Phys. Rev. **D89**, 095006 (2014).
- J.R. Batley *et al.* [NA48/2 Collab.], Phys. Lett. **B746**, 178 (2015).
- E. Abouzaid *et al.*, Phys. Rev. Lett. **107**, 201803 (2011); see also, D.G. Phillips II, “Search for the Rare Decay  $K_L \rightarrow \pi^0 \pi^0 \mu^+ \mu^-$ ,” University of Virginia thesis, May 2009.
- Y.C. Tung *et al.*, Phys. Rev. Lett. **102**, 051802 (2009).
- M. Lu and M.B. Wise, Phys. Lett. **B324**, 461 (1994); A.F. Falk, A. Lewandowski, and A.A. Petrov, Phys. Lett. **B505**, 107 (2001).
- CKMfitter Group (J. Charles *et al.*), Phys. Rev. **D84**, 033005 (2011) [[arXiv:1106.4041](#)], updated results and plots available at: <http://ckmfitter.in2p3.fr>.
- M. Bona *et al.* [UTfit Collab.] “Model-independent constraints on Delta F=2 operators and the scale of New Physics,” [arXiv:0707.0636](#), [www.utfit.org/UTfit/](http://www.utfit.org/UTfit/).
- S. Adler *et al.*, Phys. Rev. Lett. **88**, 041803 (2002).
- S. Adler *et al.*, Phys. Rev. Lett. **84**, 3768 (2000).
- S. Adler *et al.*, Phys. Lett. **B537**, 237 (2002).
- A.V. Artamonov *et al.*, Phys. Rev. Lett. **101**, 191802 (2008).
- A.V. Artamonov *et al.*, Phys. Rev. **D79**, 092004 (2009).
- G. Anelli *et al.*, CERN-SPSC-2005-013, 11 June 2005.
- J. Comfort *et al.*, FERMILAB-PROPOSAL-1021.
- G. D'Ambrosio and G. Isidori, Phys. Lett. **B530**, 108 (2002).
- J. Brod, M. Gorbahn, and E. Stamou, Phys. Rev. **D83**, 034030 (2011).
- A.J. Buras, *et al.*, JHEP **1511**, 033 (2015).
- F. Mescia and C. Smith, Phys. Rev. **D76**, 034017 (2007).
- T. Inami and C.S. Lim, Prog. Theor. Phys. **65**, 297 (1981); Erratum Prog. Theor. Phys. **65**, 172 (1981).
- G. Buchalla and A.J. Buras, Nucl. Phys. **B548**, 309 (1999); M. Misiak and J. Urban, Phys. Lett. **B451**, 161 (1999).
- A.J. Buras *et al.*, Phys. Rev. Lett. **95**, 261805 (2005); A.J. Buras *et al.*, JHEP **0611**, 002 (2006).
- J. Brod and M. Gorbahn, Phys. Rev. **D78**, 034006 (2008).
- G. Isidori, F. Mescia, and C. Smith, Nucl. Phys. **B718**, 319 (2005); A.F. Falk, A. Lewandowski, and A.A. Petrov, Phys. Lett. **B505**, 107 (2001).
- L. Littenberg and G. Valencia, Phys. Lett. **B385**, 379 (1996); C.-W. Chiang and F.J. Gilman, Phys. Rev. **D62**, 094026 (2000); C.Q. Geng, I.J. Hsu, and Y.C. Lin, Phys. Rev. **D50**, 5744 (1994).
- R. Ogata, *et al.*, Phys. Rev. **D84**, 052009 (2011).
- S. Adler, *et al.*, Phys. Rev. **D63**, 032004 (2001).
- D. Ambrose *et al.*, Phys. Rev. Lett. **84**, 1389 (2000).
- G. Valencia, Nucl. Phys. **B517**, 339 (1998).
- G. D'Ambrosio, G. Isidori, and J. Portoles, Phys. Lett. **B423**, 385 (1998).
- G. Isidori and R. Unterdorfer, JHEP **0401**, 009 (2004).
- D. Gomez-Dumm and A. Pich, Phys. Rev. Lett. **80**, 4633 (1998).
- D. Ambrose *et al.*, Phys. Rev. Lett. **81**, 4309 (1998).
- R. Aaij *et al.*, JHEP **1301**, 090 (2013).



59. L. Littenberg, Phys. Rev. **D39**, 3322 (1989).
60. G. Buchalla and G. Isidori, Phys. Lett. **B440**, 170 (1998).
61. Y. Grossman and Y. Nir, Phys. Lett. **B398**, 163 (1997).
62. J.K. Ahn *et al.*, Phys. Rev. **D81**, 072004 (2010).
63. J. Comfort *et al.*, "Proposal for  $K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$  Experiment at J-Parc," J-PARC Proposal 14 (2006).
64. K. Fuyuto, W. S. Hou, and M. Kohda, Phys. Rev. Lett. **114**, 171802 (2015).
65. K. Shiomi *et al.*, [arXiv:1411.4250](#).
66. D. Bryman *et al.*, Int. J. Mod. Phys. **A21**, 487 (2006).
67. G. D'Ambrosio *et al.*, JHEP **9808**, 004 (1998);  
C.O. Dib, I. Dunietz, and F.J. Gilman, Phys. Rev. **D39**, 2639 (1989).
68. G. Buchalla, G. D'Ambrosio, and G. Isidori, Nucl. Phys. **B672**, 387 (2003).
69. J.R. Batley *et al.*, Phys. Lett. **B576**, 43 (2003).
70. J.R. Batley *et al.*, Phys. Lett. **B599**, 197 (2004).
71. S. Friot, D. Greynat, and E. de Rafael, Phys. Lett. **B595**, 301 (2004).
72. G. Isidori, C. Smith, and R. Unterdorfer, Eur. Phys. J. **C36**, 57 (2004).
73. G. Ecker, A. Pich, and E. de Rafael, Phys. Lett. **237B**, 481 (1990);  
L. Cappiello, G. D'Ambrosio, and M. Miragliuolo, Phys. Lett. **B298**, 423 (1993);  
A. Cohen, G. Ecker, and A. Pich, Phys. Lett. **B304**, 347 (1993).
74. F. Gabbiani and G. Valencia, Phys. Rev. **D66**, 074006 (2002).
75. E. Abouzaid *et al.*, Phys. Rev. **D77**, 112004 (2008).
76. A. Lai *et al.*, Phys. Lett. **B536**, 229 (2002).
77. J. Donoghue and F. Gabbiani, Phys. Rev. **D56**, 1605 (1997).
78. E. Abouzaid *et al.*, Phys. Rev. **D76**, 052001 (2007).
79. W.M. Morse *et al.*, Phys. Rev. **D45**, 36 (1992).
80. A. Alavi-Harati *et al.*, Phys. Rev. **D64**, 012003 (2001).
81. H.B. Greenlee, Phys. Rev. **D42**, 3724 (1990).
82. A. Alavi-Harati *et al.*, Phys. Rev. Lett. **93**, 021805 (2004).
83. A. Alavi-Harati *et al.*, Phys. Rev. Lett. **84**, 5279 (2000).
84. M.V. Diwan, H. Ma, and T.L. Trueman, Phys. Rev. **D65**, 054020 (2002).
85. F. Mescia, C. Smith, and S. Trine, JHEP **0608**, 088 (2006).
86. R. Appel *et al.*, Phys. Rev. Lett. **83**, 4482 (1999);  
J.R. Batley *et al.*, Phys. Lett. **B677**, 246 (2009).
87. S.C. Adler *et al.*, Phys. Rev. Lett. **79**, 4756 (1997);  
R. Appel *et al.*, Phys. Rev. Lett. **84**, 2580 (2000);  
H.K. Park *et al.*, Phys. Rev. Lett. **88**, 111801 (2002).
88. R. Fantechi, Pos HQL **2012**, 014 (2012).
89. L. Cappiello, *et al.*, Eur. Phys. J. **C72**, 1872 (2012) [Eur. Phys. J. **C72**, 2208 (2012)] [[arXiv:1112.5184](#)].
90. G. Ecker, A. Pich, and E. de Rafael, Nucl. Phys. **B303**, 665 (1988).
91. G. D'Ambrosio and J. Portoles, Phys. Lett. **B386**, 403 (1996) [Phys. Lett. **B389**, 770 (1996)] [Erratum-ibid. **B395**, 390 (1997)] [[hep-ph/9606213](#)].
92. P. Kitching *et al.* [E787 Collab.], Phys. Rev. Lett. **79**, 4079 (1997) [[hep-ex/9708011](#)].

93. J.R. Batley *et al.*, Phys. Lett. **B730**, 141 (2014).
94. C. Lazzeroni *et al.*, Phys. Lett. **B732C**, 65 (2014).
95. A. Alavi-Harati *et al.*, Phys. Rev. Lett. **87**, 071801 (2001).
96. A. Abouzaid *et al.*, Phys. Rev. Lett. **99**, 051804 (2007).
97. J.R. LaDue "Understanding Dalitz Decays of the  $K_L$  in particular the decays of  $K_L \rightarrow e^+ e^- \gamma$  and  $K_L \rightarrow e^+ e^- e^+ e^-$ " University of Colorado Thesis, May 2003. The preliminary result for  $K_L \rightarrow e^+ e^- \gamma$  in this thesis has been superseded by the final result in [96].
98. A. Alavi-Harati *et al.*, Phys. Rev. Lett. **86**, 5425 (2001).
99. V. Fanti *et al.*, Phys. Lett. **B458**, 458 (1999).

 **$K^+$  DECAY MODES** $K^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Leptonic and semileptonic modes</b>		
$\Gamma_1$ $e^+ \nu_e$	( $1.582 \pm 0.007$ ) $\times 10^{-5}$	
$\Gamma_2$ $\mu^+ \nu_\mu$	( $63.56 \pm 0.11$ ) %	S=1.2
$\Gamma_3$ $\pi^0 e^+ \nu_e$	( $5.07 \pm 0.04$ ) %	S=2.1
Called $K_{23}^+$ .		
$\Gamma_4$ $\pi^0 \mu^+ \nu_\mu$	( $3.352 \pm 0.033$ ) %	S=1.9
Called $K_{\mu 3}^+$ .		
$\Gamma_5$ $\pi^0 \pi^0 e^+ \nu_e$	( $2.55 \pm 0.04$ ) $\times 10^{-5}$	S=1.1
$\Gamma_6$ $\pi^+ \pi^- e^+ \nu_e$	( $4.247 \pm 0.024$ ) $\times 10^{-5}$	
$\Gamma_7$ $\pi^+ \pi^- \mu^+ \nu_\mu$	( $1.4 \pm 0.9$ ) $\times 10^{-5}$	
$\Gamma_8$ $\pi^0 \pi^0 \pi^0 e^+ \nu_e$	< 3.5 $\times 10^{-6}$	CL=90%
<b>Hadronic modes</b>		
$\Gamma_9$ $\pi^+ \pi^0$	( $20.67 \pm 0.08$ ) %	S=1.2
$\Gamma_{10}$ $\pi^+ \pi^0 \pi^0$	( $1.760 \pm 0.023$ ) %	S=1.1
$\Gamma_{11}$ $\pi^+ \pi^+ \pi^-$	( $5.583 \pm 0.024$ ) %	
<b>Leptonic and semileptonic modes with photons</b>		
$\Gamma_{12}$ $\mu^+ \nu_\mu \gamma$	[a,b] ( $6.2 \pm 0.8$ ) $\times 10^{-3}$	
$\Gamma_{13}$ $\mu^+ \nu_\mu \gamma$ (SD <sup>+</sup> )	[c,d] ( $1.33 \pm 0.22$ ) $\times 10^{-5}$	
$\Gamma_{14}$ $\mu^+ \nu_\mu \gamma$ (SD <sup>+</sup> INT)	[c,d] < 2.7 $\times 10^{-5}$	CL=90%
$\Gamma_{15}$ $\mu^+ \nu_\mu \gamma$ (SD <sup>-</sup> + SD <sup>-</sup> INT)	[c,d] < 2.6 $\times 10^{-4}$	CL=90%
$\Gamma_{16}$ $e^+ \nu_e \gamma$	( $9.4 \pm 0.4$ ) $\times 10^{-6}$	
$\Gamma_{17}$ $\pi^0 e^+ \nu_e \gamma$	[a,b] ( $2.56 \pm 0.16$ ) $\times 10^{-4}$	
$\Gamma_{18}$ $\pi^0 e^+ \nu_e \gamma$ (SD)	[c,d] < 5.3 $\times 10^{-5}$	CL=90%
$\Gamma_{19}$ $\pi^0 \mu^+ \nu_\mu \gamma$	[a,b] ( $1.25 \pm 0.25$ ) $\times 10^{-5}$	
$\Gamma_{20}$ $\pi^0 \pi^0 e^+ \nu_e \gamma$	< 5 $\times 10^{-6}$	CL=90%
<b>Hadronic modes with photons or <math>\ell\bar{\ell}</math> pairs</b>		
$\Gamma_{21}$ $\pi^+ \pi^0 \gamma$ (INT)	( - 4.2 $\pm 0.9$ ) $\times 10^{-6}$	
$\Gamma_{22}$ $\pi^+ \pi^0 \gamma$ (DE)	[a,e] ( $6.0 \pm 0.4$ ) $\times 10^{-6}$	
$\Gamma_{23}$ $\pi^+ \pi^0 \pi^0 \gamma$	[a,b] ( $7.6 \pm 6.0$ ) $\times 10^{-6}$	
$\Gamma_{24}$ $\pi^+ \pi^+ \pi^- \gamma$	[a,b] ( $1.04 \pm 0.31$ ) $\times 10^{-4}$	
$\Gamma_{25}$ $\pi^+ \gamma \gamma$	[a] ( $1.01 \pm 0.06$ ) $\times 10^{-6}$	
$\Gamma_{26}$ $\pi^+ 3\gamma$	[a] < 1.0 $\times 10^{-4}$	CL=90%
$\Gamma_{27}$ $\pi^+ e^+ e^- \gamma$	( $1.19 \pm 0.13$ ) $\times 10^{-8}$	
<b>Leptonic modes with <math>\ell\bar{\ell}</math> pairs</b>		
$\Gamma_{28}$ $e^+ \nu_e \nu \bar{\nu}$	< 6 $\times 10^{-5}$	CL=90%
$\Gamma_{29}$ $\mu^+ \nu_\mu \nu \bar{\nu}$	< 6.0 $\times 10^{-6}$	CL=90%
$\Gamma_{30}$ $e^+ \nu_e e^+ e^-$	( $2.48 \pm 0.20$ ) $\times 10^{-8}$	
$\Gamma_{31}$ $\mu^+ \nu_\mu e^+ e^-$	( $7.06 \pm 0.31$ ) $\times 10^{-8}$	
$\Gamma_{32}$ $e^+ \nu_e \mu^+ \mu^-$	( $1.7 \pm 0.5$ ) $\times 10^{-8}$	
$\Gamma_{33}$ $\mu^+ \nu_\mu \mu^+ \mu^-$	< 4.1 $\times 10^{-7}$	CL=90%
<b>Lepton family number (LF), Lepton number (L), <math>\Delta S = \Delta Q</math> (SQ) violating modes, or <math>\Delta S = 1</math> weak neutral current (S1) modes</b>		
$\Gamma_{34}$ $\pi^+ \pi^+ e^- \bar{\nu}_e$	SQ < 1.3 $\times 10^{-8}$	CL=90%
$\Gamma_{35}$ $\pi^+ \pi^+ \mu^- \bar{\nu}_\mu$	SQ < 3.0 $\times 10^{-6}$	CL=95%
$\Gamma_{36}$ $\pi^+ e^+ e^-$	S1 ( $3.00 \pm 0.09$ ) $\times 10^{-7}$	
$\Gamma_{37}$ $\pi^+ \mu^+ \mu^-$	S1 ( $9.4 \pm 0.6$ ) $\times 10^{-8}$	S=2.6

# Meson Particle Listings

$K^\pm$

$\Gamma_{38}$	$\pi^+ \nu \overline{\nu}$	$SI$	(	1.7	$\pm 1.1$	) $\times 10^{-10}$	
$\Gamma_{39}$	$\pi^+ \pi^0 \nu \overline{\nu}$	$SI$	<	4.3		$\times 10^{-5}$	CL=90%
$\Gamma_{40}$	$\mu^- \nu e^+ e^+$	$LF$	<	2.1		$\times 10^{-8}$	CL=90%
$\Gamma_{41}$	$\mu^+ \nu_e$	$LF$	[f]	<	4	$\times 10^{-3}$	CL=90%
$\Gamma_{42}$	$\pi^+ \mu^+ e^-$	$LF$	<	1.3		$\times 10^{-11}$	CL=90%
$\Gamma_{43}$	$\pi^+ \mu^- e^+$	$LF$	<	5.2		$\times 10^{-10}$	CL=90%
$\Gamma_{44}$	$\pi^- \mu^+ e^+$	$L$	<	5.0		$\times 10^{-10}$	CL=90%
$\Gamma_{45}$	$\pi^- e^+ e^+$	$L$	<	6.4		$\times 10^{-10}$	CL=90%
$\Gamma_{46}$	$\pi^- \mu^+ \mu^+$	$L$	[f]	<	1.1	$\times 10^{-9}$	CL=90%
$\Gamma_{47}$	$\mu^+ \overline{\nu}_e$	$L$	[f]	<	3.3	$\times 10^{-3}$	CL=90%
$\Gamma_{48}$	$\pi^0 e^+ \overline{\nu}_e$	$L$	<	3		$\times 10^{-3}$	CL=90%
$\Gamma_{49}$	$\pi^+ \gamma$	[g]	<	2.3		$\times 10^{-9}$	CL=90%

- [a] See the Particle Listings below for the energy limits used in this measurement.
- [b] Most of this radiative mode, the low-momentum  $\gamma$  part, is also included in the parent mode listed without  $\gamma$ 's.
- [c] Structure-dependent part.
- [d] See the "Note on  $\pi^\pm \rightarrow \ell^\pm \nu \gamma$  and  $K^\pm \rightarrow \ell^\pm \nu \gamma$  Form Factors" in the  $\pi^\pm$  Particle Listings for definitions and details.
- [e] Direct-emission branching fraction.
- [f] Derived from an analysis of neutrino-oscillation experiments.
- [g] Violates angular-momentum conservation.

## CONSTRAINED FIT INFORMATION

An overall fit to the mean life, a decay rate, and 15 branching ratios uses 35 measurements and one constraint to determine 8 parameters. The overall fit has a  $\chi^2 = 53.4$  for 28 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	−66						
$x_4$	−64	90					
$x_5$	−12	−5	−5				
$x_9$	−67	0	−1	−6			
$x_{10}$	−13	−6	−5	91	−6		
$x_{11}$	−14	−6	−6	2	−7	2	
$\Gamma$	3	1	1	0	2	0	−24
	$x_2$	$x_3$	$x_4$	$x_5$	$x_9$	$x_{10}$	$x_{11}$

Mode	Rate ( $10^8 \text{ s}^{-1}$ )	Scale factor
$\Gamma_2$ $\mu^+ \nu_\mu$	0.5134 $\pm$ 0.0012	1.5
$\Gamma_3$ $\pi^0 e^+ \nu_e$ Called $K_{e3}^+$ .	0.0410 $\pm$ 0.0004	2.1
$\Gamma_4$ $\pi^0 \mu^+ \nu_\mu$ Called $K_{\mu3}^+$ .	0.02707 $\pm$ 0.00027	1.9
$\Gamma_5$ $\pi^0 \pi^0 e^+ \nu_e$	(2.059 $\pm$ 0.029) $\times 10^{-5}$	1.1
$\Gamma_9$ $\pi^+ \pi^0$	0.1670 $\pm$ 0.0007	1.3
$\Gamma_{10}$ $\pi^+ \pi^0 \pi^0$	0.01421 $\pm$ 0.00018	1.1
$\Gamma_{11}$ $\pi^+ \pi^+ \pi^-$	0.04510 $\pm$ 0.00019	

## $K^\pm$ DECAY RATES

$\Gamma(\mu^+\nu_\mu)$	$\Gamma_2$		
<u>VALUE (<math>10^6 \text{ s}^{-1}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>
<b>51.34<math>\pm</math>0.12 OUR FIT</b>	Error includes scale factor of 1.5.		
• • • We do not use the following data for averages, fits, limits, etc. • • •			
51.2 $\pm$ 0.8	FORD	67	CNTR $\pm$

$\Gamma(\pi^+\pi^+\pi^-)$				$\Gamma_{11}$
VALUE ( $10^6 \text{ s}^{-1}$ )	EVTS	DOCUMENT ID	TECN	CHG
4.510 $\pm$ 0.019 OUR FIT				
4.511 $\pm$ 0.024		1 FORD	70	ASPK
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.529 $\pm$ 0.032	3.2M	1 FORD	70	ASPK
4.496 $\pm$ 0.030		1 FORD	67	CNTR $\pm$
1 First FORD 70 value is second FORD 70 combined with FORD 67.				

## $K^+$ BRANCHING RATIOS

### Leptonic and semileptonic modes

$\Gamma(e^+\nu_e)/\Gamma(\mu^+\nu_\mu)$				$\Gamma_1/\Gamma_2$
See the note on “Decay Constants of Charged Pseudoscalar Mesons” in the $D_s^+$ Listings.				
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>
<b>2.488±0.009 OUR AVERAGE</b>				
2.488±0.007±0.007	150k	<sup>1</sup> LAZZERONI	13	NA62 ±
2.493±0.025±0.019	13.8K	<sup>2</sup> AMBROSINO	09E	KLOE ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.487±0.011±0.007	60k	<sup>3</sup> LAZZERONI	11	NA62 +
2.51 ±0.15	404	HEINTZE	76	SPEC +
2.37 ±0.17	534	HEARD	75B	SPEC +
2.42 ±0.42	112	CLARK	72	OSP K +

- 1 LAZZERONI 13 uses full data sample collected from 2007 to 2008. This ratio is defined to be fully inclusive, including internal-bremsstrahlung.
- 2 The ratio is defined to include internal-bremsstrahlung, ignoring direct-emission contributions. AMBROSINO 09E determined the ratio from the measurement of  $\Gamma(K \rightarrow e \nu(\gamma))$ ,  $E_\gamma < 10 \text{ MeV}$  /  $\Gamma(K \rightarrow \mu \nu(\gamma))$ . 89.8% of  $K \rightarrow e \nu(\gamma)$  events had  $E_\gamma < 10 \text{ MeV}$ .
- 3 This ratio is defined to be fully inclusive, including internal-bremsstrahlung.

$\Gamma(\mu^+ \nu_\mu)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$				
See the note on "Decay Constants of Charged Pseudoscalar Mesons" in the $D_s^+$ Listings.					
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>63.56<math>\pm</math>0.11 OUR FIT</b>	Error includes scale factor of 1.2.				
<b>63.60<math>\pm</math>0.16 OUR AVERAGE</b>					
63.66 $\pm$ 0.09 $\pm$ 0.15	865k	<sup>1</sup> AMBROSINO	06A	KLOE	+
63.24 $\pm$ 0.44	62k	CHIANG	72	OSP K	+
1.84 GeV/c $K^+$					

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>5.07 <math>\pm</math> 0.04 OUR FIT</b>	Error includes scale factor of 2.1.				
<b>4.94 <math>\pm</math> 0.05 OUR AVERAGE</b>					
4.965 $\pm$ 0.038 $\pm$ 0.037		<sup>1</sup> AMBROSINO	08A	KLOE	$\pm$
4.86 $\pm$ 0.10	3516	CHIANG	72	OSP K	$+$ 1.84 GeV/c $K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.7 $\pm$ 0.3	429	SHAKLEE	64	HLBC	$+$
5.0 $\pm$ 0.5		ROE	61	HLBC	$+$

- 1 Depends on  $K^+$  lifetime  $\tau$ . AMBROSINO 08A uses PDG 06 value of  $\tau = (1.2385 \pm 0.0024) \times 10^{-8} \text{ sec}$ . The correlation between  $K_{e3}^+$  and  $K_{\mu3}^+$  branching fraction measurements is 62.7%.

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$					$\Gamma_3/\Gamma_2$	
VALUE		EVTS	DOCUMENT ID	TECN	CHG	
$0.0798 \pm 0.0008$	OUR FIT	Error includes scale factor of 1.9.				
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.069 \pm 0.006$		350	ZELLER	69	ASPK +	
$0.0775 \pm 0.0033$		960	BOTTERILL	68c	ASPK +	
$0.069 \pm 0.006$		561	GARLAND	68	OSP K +	
$0.0791 \pm 0.0054$		295	<sup>1</sup> AUERBACH	67	OSP K +	

- 1 AUERBACH 67 changed from 0.0797  $\pm$  0.0054. See comment with ratio  $\Gamma(\pi^0 \mu^+ \nu_\mu) / \Gamma(\mu^+ \nu_\mu)$ . The value 0.0785  $\pm$  0.0025 given in AUERBACH 67 is an average of AUERBACH 67  $\Gamma(\pi^0 e^+ \nu_e) / \Gamma(\mu^+ \nu_\mu)$  and CESTER 66  $\Gamma(\pi^0 e^+ \nu_e) / [\Gamma(\mu^+ \nu_\mu) + \Gamma(\pi^+ \pi^0)]$ .

$\Gamma(\pi^0 e^+ \nu_e) / [\Gamma(\mu^+ \nu_\mu) + \Gamma(\pi^+ \pi^0)]$				$\Gamma_3/(\Gamma_2+\Gamma_9)$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>6.02±0.06 OUR FIT</b>	Error includes scale factor of 2.1.			
<b>6.02±0.15 OUR AVERAGE</b>				
6.16±0.22	5110	ESCHSTRUTH	68	OSP K +
5.89±0.21	1679	CESTER	66	OSP K +
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.92±0.65		<sup>1</sup> WEISSENBERG...	76	SPEC +
<sup>1</sup> Value calculated from WEISSENBERG 76 ( $\pi^0 e \nu$ ), ( $\mu \nu$ ), and ( $\pi \pi^0$ ) values to eliminate dependence on our 1974 ( $\pi 2\pi^0$ ) and ( $\pi^+ \pi^-$ ) fractions.				

$\Gamma(\pi^0 e^+ \nu_e) / [\Gamma(\pi^0 \mu^+ \nu_\mu) + \Gamma(\pi^+ \pi^0) + \Gamma(\pi^+ \pi^0 \pi^0)]$				$\Gamma_3 / (\Gamma_4 + \Gamma_9 + \Gamma_{10})$
VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.1967 ± 0.0016 OUR FIT</b>	Error includes scale factor of 2.5.			
<b>0.1962 ± 0.0008 ± 0.0035</b>	71k	SHER	03 B865	+

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\pi^+ \pi^0)$					$\Gamma_3/\Gamma_9$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.2454 ± 0.0023 OUR FIT</b>		Error includes scale factor of 2.6.			
<b>0.2467 ± 0.0011 OUR AVERAGE</b>		Error includes scale factor of 1.1.			
0.2423 ± 0.0015 ± 0.0037	31k	UVAROV	14	ISTR	— ISTR+
0.2470 ± 0.0009 ± 0.0004	87k	BATLEY	07A	NA48	±
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.221 ± 0.012	786	<sup>1</sup> LUCAS	73B	HBC	— Dalitz pairs only
<sup>1</sup> LUCAS 73B gives $N(K_{e3}) = 786 \pm 3.1\%$ , $N(2\pi) = 3564 \pm 3.1\%$ . We use these values to obtain quoted result.					

See key on page 601

## Meson Particle Listings

 $K^\pm$  $\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\pi^+ \pi^+ \pi^-)$  $\Gamma_3/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.908±0.009 OUR FIT</b>	Error	includes scale factor of 1.6.		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.867±0.027	2768	BARMIN	87	XEBC +
0.856±0.040	2827	BRAUN	75	HLBC +
0.850±0.019	4385	<sup>1</sup> HAIDT	71	HLBC +
0.846±0.021	4385	<sup>1</sup> EICHTE	68	HLBC +
0.94 ± 0.09	854	BELLOTTI	67B	HLBC
0.90 ± 0.06	230	BORREANI	64	HBC +

<sup>1</sup>HAIDT 71 is a reanalysis of EICHTE 68. Not included in average because of large discrepancy in  $\Gamma(\pi^0 \mu^+ \nu)/\Gamma(\pi^0 e^+ \nu)$  with more precise results.

 $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma_{\text{total}}$  $\Gamma_4/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>3.352±0.033 OUR FIT</b>	Error	includes scale factor of 1.9.			
<b>3.24 ± 0.04 OUR AVERAGE</b>					
3.233±0.029±0.026		<sup>1</sup> AMBROSINO	08A	KLOE	±
3.33 ± 0.16	2345	CHIANG	72	OSPK	+ 1.84 GeV/c $K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.8 ± 0.4		<sup>2</sup> TAYLOR	59	EMUL	+

<sup>1</sup>Depends on  $K^+$  lifetime  $\tau$ . AMBROSINO 08A uses PDG 06 value of  $\tau = (1.2385 \pm 0.0024) \times 10^{-8}$  sec. The correlation between  $K_{e3}^+$  and  $K_{\mu 3}^+$  branching fraction measurements is 62.7%.

<sup>2</sup>Earlier experiments not averaged.

 $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\mu^+ \nu_\mu)$  $\Gamma_4/\Gamma_2$ 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.0527±0.0006 OUR FIT</b>	Error	includes scale factor of 1.8.		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.054 ± 0.009	240	ZELLER	69	ASPK +
0.0480±0.0037	424	<sup>1</sup> GARLAND	68	OSPK +
0.0486±0.0040	307	<sup>2</sup> AUERBACH	67	OSPK +

<sup>1</sup>GARLAND 68 changed from 0.055 ± 0.004 in agreement with  $\mu$ -spectrum calculation of GAILLARD 70 appendix B. L.G.Pondrom, (private communication 73).

<sup>2</sup>AUERBACH 67 changed from 0.0602 ± 0.0046 by erratum which brings the  $\mu$ -spectrum calculation into agreement with GAILLARD 70 appendix B.

 $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^+ e^+ \nu_e)$  $\Gamma_4/\Gamma_3$ 

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.6608±0.0029 OUR FIT</b>	Error	includes scale factor of 1.1.			
<b>0.6618±0.0027 OUR AVERAGE</b>					
0.663 ± 0.003 ± 0.001	77k	BATLEY	07A	NA48	±
0.671 ± 0.007 ± 0.008	24k	HORIE	01	SPEC	
0.670 ± 0.014		<sup>1</sup> HEINTZE	77	SPEC	+
0.667 ± 0.017	5601	BOTTERILL	68B	ASPK	+
• • • We use the following data for averages but not for fits. • • •					
0.6511±0.0064		<sup>2</sup> AMBROSINO	08A	KLOE	±
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.608 ± 0.014	1585	<sup>3</sup> BRAUN	75	HLBC	+
0.705 ± 0.063	554	<sup>4</sup> LUCAS	73B	HBC	− Dalitz pairs only
0.698 ± 0.025	3480	<sup>5</sup> CHIANG	72	OSPK	+ 1.84 GeV/c $K^+$
0.596 ± 0.025		<sup>6</sup> HAIDT	71	HLBC	+
0.604 ± 0.022	1398	<sup>6</sup> EICHTE	68	HLBC	
0.703 ± 0.056	1509	CALLAHAN	66B	HLBC	

<sup>1</sup>HEINTZE 77 value from fit to  $\lambda_0$ . Assumes  $\mu$ -e universality.

<sup>2</sup>Not used in the fit. This result enters the fit via correlation of  $K_{e3}^+$  and  $K_{\mu 3}^+$  branching fraction measurements of AMBROSINO 08A.

<sup>3</sup>BRAUN 75 value is from form factor fit. Assumes  $\mu$ -e universality.

<sup>4</sup>LUCAS 73B gives  $N(K_{\mu 3}) = 554 \pm 7.6\%$ ,  $N(K_{e3}) = 786 \pm 3.1\%$ . We divide.

<sup>5</sup>CHIANG 72  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^0 e^+ \nu_e)$  is statistically independent of CHIANG 72  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma_{\text{total}}$  and  $\Gamma(\pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$ .

<sup>6</sup>HAIDT 71 is a reanalysis of EICHTE 68. Not included in average because of large discrepancy with more precise results.

 $[\Gamma(\pi^0 \mu^+ \nu_\mu) + \Gamma(\pi^+ \pi^0)]/\Gamma_{\text{total}}$  $(\Gamma_4 + \Gamma_9)/\Gamma$ 

We combine these two modes for experiments measuring them in xenon bubble chamber because of difficulties of separating them there.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>24.02±0.08 OUR FIT</b>	Error	includes scale factor of 1.2.		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
25.4 ± 0.9	886	SHAKLEE	64	HLBC +
23.4 ± 1.1		ROE	61	HLBC +

 $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^0)$  $\Gamma_4/\Gamma_9$ 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.1637±0.0006±0.0003</b>	77k	BATLEY	07A	NA48 ±

 $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$  $\Gamma_4/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.600±0.007 OUR FIT</b>	Error	includes scale factor of 1.6.			
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.503±0.019	1505	<sup>1</sup> HAIDT	71	HLBC	+
0.510±0.017	1505	<sup>1</sup> EICHTE	68	HLBC	+
0.63 ± 0.07	2845	<sup>2</sup> BISI	65B	BC	+ HBC+HLBC

<sup>1</sup>HAIDT 71 is a reanalysis of EICHTE 68. Not included in average because of large discrepancy in  $\Gamma(\pi^0 \mu^+ \nu)/\Gamma(\pi^0 e^+ \nu)$  with more precise results.

<sup>2</sup>Error enlarged for background problems. See GAILLARD 70.

 $\Gamma(\pi^0 \pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$  $\Gamma_5/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>2.55±0.04 OUR FIT</b>	Error	includes scale factor of 1.1.		
<b>2.54±0.89</b>	10	BARMIN	88B	HLBC +

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>1.449±0.008 OUR FIT</b>				
<b>1.449±0.006±0.006</b>	65.2k	<sup>1</sup> BATLEY	14A	NA48 ±

<sup>1</sup>Data collected in 2003–2004. This leads to the scalar form factor  $(1 + \delta_{EM}) f_3 = 6.079 \pm 0.012 \pm 0.027 \pm 0.046$  where the last error is due to the normalizing decay mode uncertainty.

 $\Gamma(\pi^0 \pi^0 e^+ \nu_e)/\Gamma(\pi^0 e^+ \nu_e)$  $\Gamma_5/\Gamma_3$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>5.03±0.09 OUR FIT</b>	Error	includes scale factor of 1.2.		
<b>4.1 <math>\pm 1.0</math> <math>-0.7</math> OUR AVERAGE</b>				
4.2 $\pm 1.0$ $-0.9$	25	BOLOTOV	86B	CALO −
3.8 $\pm 5.0$ $-1.2$	2	LJUNG	73	HLBC +

 $\Gamma(\pi^+ \pi^- e^+ \nu_e)/\Gamma(\pi^+ \pi^+ \pi^-)$  $\Gamma_6/\Gamma_{11}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>7.606±0.029 OUR AVERAGE</b>				
7.615±0.008±0.028	1.1M	<sup>1</sup> BATLEY	12	NA48 ±
7.35 ± 0.01 ± 0.19	388k	<sup>2</sup> PISLAK	01	B865
7.21 ± 0.32	30k	ROSSELET	77	SPEC +
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7.36 ± 0.68	500	BOURQUIN	71	ASPK
7.0 ± 0.9	106	SCHWEIN...	71	HLBC +
5.83 ± 0.63	269	ELY	69	HLBC +

<sup>1</sup>BATLEY 12 uses data collected in 2003–2004. The result is inclusive of  $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu_\gamma$  decays. Using PDG 12 value for  $\Gamma(\pi^+ \pi^- \pi^+)/\Gamma = (5.59 \pm 0.04) \times 10^{-2}$ . BATLEY 12 obtains  $B(\pi^+ \pi^- e \nu) = (4.257 \pm 0.004 \pm 0.035) \times 10^{-5}$  where the syst. error is dominated by the error on the normalization mode.

<sup>2</sup>PISLAK 01 reports  $\Gamma(\pi^+ \pi^- e^+ \nu_e)/\Gamma_{\text{total}} = (4.109 \pm 0.008 \pm 0.110) \times 10^{-5}$  using the PDG 00 value  $\Gamma(\pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}} = (5.59 \pm 0.05) \times 10^{-2}$ . We divide by the PDG value and unfold its error from the systematic error. PISLAK 03 and PISLAK 10A give additional details on the branching ratio measurement and give improved errors on the S-wave  $\pi$ - $\pi$  scattering length:  $a_0^0 = 0.235 \pm 0.013$  and  $a_0^0 = -0.0410 \pm 0.0027$ .

 $\Gamma(\pi^+ \pi^- \mu^+ \nu_\mu)/\Gamma_{\text{total}}$  $\Gamma_7/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;3.5</b>	90	BOLOTOV	88	SPEC −
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.77 $\pm 0.54$ $-0.50$	1	CLINE	65	FBC +

 $\Gamma(\pi^+ \pi^- \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$  $\Gamma_7/\Gamma_{11}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>2.57±1.55</b>	7	BISI	67	DBC +
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 2.5	1	GREINER	64	EMUL +

 $\Gamma(\pi^0 \pi^0 \pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$  $\Gamma_8/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;3.5</b>	90	0	BOLOTOV	88	SPEC −
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<9	90	0	BARMIN	92	XEBC +

## Hadronic modes

 $\Gamma(\pi^+ \pi^0)/\Gamma_{\text{total}}$  $\Gamma_9/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>20.67±0.08 OUR FIT</b>	Error	includes scale factor of 1.2.			
<b>20.70±0.16 OUR AVERAGE</b>	Error	includes scale factor of 1.8.			
20.65±0.05±0.08	1.4M	<sup>1</sup> AMBROSINO	08E	KLOE	+ $\phi \rightarrow K^+ K^-$
21.18±0.28	16k	CHIANG	72	OSPK	+ 1.84 GeV/c $K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
21.0 ± 0.6		CALLAHAN	65	HLBC	See $\Gamma_9/\Gamma_{11}$

<sup>1</sup>Fully inclusive of final-state radiation. The branching ratio is evaluated using  $K^+$  lifetime,  $\tau = 12.385$  ns.

# Meson Particle Listings

$K^\pm$

$\Gamma(\pi^+\pi^0)/\Gamma(\pi^+\pi^+\pi^-)$					$\Gamma_9/\Gamma_{11}$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	
<b><math>0.3252\pm0.022</math> OUR FIT</b>	Error includes scale factor of 1.1.				
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
3.96 ±0.15	1045	CALLAHAN	66	FBC	+

$\Gamma(\pi^+\pi^0)/\Gamma(\mu^+\nu_\mu)$					$\Gamma_9/\Gamma_2$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.3252\pm0.0016</math> OUR FIT</b>	Error includes scale factor of 1.2.				
<b><math>0.3325\pm0.0032</math> OUR AVERAGE</b>					
0.3329±0.0047±0.0010	45k	USHER	92	SPEC	+ $p\overline{p}$ at rest
0.3355±0.0057		<sup>1</sup> WEISSENBE...	76	SPEC	+
0.3277±0.0065	4517	<sup>2</sup> AUERBACH	67	OSPK	+
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
0.328 ±0.005	25k	<sup>1</sup> WEISSENBE...	74	STRC	+
0.305 ±0.018	1600	ZELLER	69	ASPK	+
<sup>1</sup> WEISSENBERG 76 revises WEISSENBERG 74.					
<sup>2</sup> AUERBACH 67 changed from 0.3253 ± 0.0065. See comment with ratio $\Gamma(\pi^0\mu^+\nu_\mu)/\Gamma(\mu^+\nu_\mu)$ .					

$\Gamma(\pi^+\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{10}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1.760\pm0.023</math> OUR FIT</b>	Error includes scale factor of 1.1.				
<b><math>1.775\pm0.028</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.				
1.763±0.013±0.022		ALOISIO	04A	KLOE	±
1.84 ±0.06	1307	CHIANG	72	OSPK	+ 1.84 GeV/c $K^+$
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
1.53 ±0.11	198	<sup>1</sup> PANDOULAS	70	EMUL	+
1.8 ±0.2	108	SHAKLEE	64	HLBC	+
1.7 ±0.2		ROE	61	HLBC	+
1.5 ±0.2		<sup>2</sup> TAYLOR	59	EMUL	+
<sup>1</sup> Includes events of TAYLOR 59.					
<sup>2</sup> Earlier experiments not averaged.					

$\Gamma(\pi^+\pi^0\pi^0)/\Gamma(\pi^+\pi^0)$					$\Gamma_{10}/\Gamma_9$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.0851\pm0.0012</math> OUR FIT</b>	Error includes scale factor of 1.1.				
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
0.081 ±0.005	574	<sup>1</sup> LUCAS	73B	HBC	— Dalitz pairs only
<sup>1</sup> LUCAS 73B gives $N(\pi^2\pi^0) = 574 \pm 5.9\%$ , $N(2\pi) = 3564 \pm 3.1\%$ . We quote $0.5N(\pi^2\pi^0)/N(2\pi)$ where 0.5 is because only Dalitz pair $\pi^0$ 's were used.					

$\Gamma(\pi^+\pi^0\pi^0)/\Gamma(\pi^+\pi^+\pi^-)$					$\Gamma_{10}/\Gamma_{11}$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.315\pm0.004</math> OUR FIT</b>	Error includes scale factor of 1.1.				
<b><math>0.303\pm0.009</math></b>	2027	BISI	65	BC	+ HBC+HLBC
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
0.393±0.099	17	YOUNG	65	EMUL	+

$\Gamma(\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>5.583\pm0.024</math> OUR FIT</b>	Error includes scale factor of 1.3.				
<b><math>5.565\pm0.031\pm0.025</math></b>	68K	<sup>1</sup> BABUSCI	14B	KLOE	+
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
5.56 ±0.20	2330	<sup>2</sup> CHIANG	72	OSPK	+ 1.84 GeV/c $K^+$
5.34 ±0.21	693	<sup>3</sup> PANDOULAS	70	EMUL	+
5.71 ±0.15		DEMARCO	65	HBC	
6.0 ±0.4	44	YOUNG	65	EMUL	+
5.54 ±0.12	2332	CALLAHAN	64	HLBC	+
5.1 ±0.2	540	SHAKLEE	64	HLBC	+
5.7 ±0.3		ROE	61	HLBC	+

- <sup>1</sup>Inclusive of final-state radiation. Result obtained from averaging two branching ratios: one from a sample with  $K^- \rightarrow \mu\nu(\gamma)$  tagging and another with  $K^- \rightarrow \pi^-\pi^0(\gamma)$  tagging.

<sup>2</sup>Value is not independent of CHIANG 72  $\Gamma(\mu^+\nu_\mu)/\Gamma_{\text{total}}$ ,  $\Gamma(\pi^+\pi^0)/\Gamma_{\text{total}}$ ,  $\Gamma(\pi^+\pi^0\pi^0)/\Gamma_{\text{total}}$ ,  $\Gamma(\pi^0\mu^+\nu_\mu)/\Gamma_{\text{total}}$ , and  $\Gamma(\pi^0e^+\nu_e)/\Gamma_{\text{total}}$ .

<sup>3</sup>Includes events of TAYLOR 59.

## Leptonic and semileptonic modes with photons

$\Gamma(\mu^+\nu_\mu\gamma)/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>6.2\pm0.8</math> OUR AVERAGE</b>					
6.6±1.5		<sup>1,2</sup> DEMIDOV	90	XEBC	$P(\mu) < 231.5$ MeV/c
6.0±0.9		BARMIN	88	HLBC	+ $P(\mu) < 231.5$ MeV/c
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
3.5±0.8		<sup>2,3</sup> DEMIDOV	90	XEBC	$E(\gamma) > 20$ MeV
3.2±0.5	57	<sup>4</sup> BARMIN	88	HLBC	+ $E(\gamma) > 20$ MeV
5.4±0.3		<sup>5</sup> AKIBA	85	SPEC	$P(\mu) < 231.5$ MeV/c

- <sup>1</sup> $P(\mu)$  cut given in DEMIDOV 90 paper, 235.1 MeV/c, is a misprint according to authors (private communication).

<sup>2</sup>DEMIDOV 90 quotes only inner bremsstrahlung (IB) part.

<sup>3</sup>Not independent of above DEMIDOV 90 value. Cuts differ.

<sup>4</sup>Not independent of above BARMIN 88 value. Cuts differ.

<sup>5</sup>Assumes  $\mu$ -e universality and uses constraints from  $K \rightarrow e\nu\gamma$ .

$\Gamma(\mu^+\nu_\mu\gamma(\text{SD}^+))/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
Structure-dependent part with $+\gamma$ helicity ( $\text{SD}^+$ term). See the “Note on $\pi^\pm \rightarrow \ell^\pm\nu\gamma$ and $K^\pm \rightarrow \ell^\pm\nu\gamma$ Form Factors” in the $\pi^\pm$ section of the Particle Data Listings above.					
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	
<b><math>1.33\pm0.12\pm0.18</math></b>	2588	<sup>1</sup> ADLER	00B	B787	
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
<3.0	90	AKIBA	85	SPEC	

- <sup>1</sup>ADLER 00B obtains the branching ratio by extrapolating the measurement in the kinematic region  $E_\mu > 137$  MeV,  $E_\gamma > 90$  MeV to the full  $\text{SD}^+$  phase-space. Also reports  $|F_V + F_A| = 0.165 \pm 0.007 \pm 0.011$  and  $-0.04 < F_V - F_A < 0.24$  at 90% CL.

$\Gamma(\mu^+\nu_\mu\gamma(\text{SD}^+\text{INT}))/\Gamma_{\text{total}}$					$\Gamma_{14}/\Gamma$
Interference term between internal Bremsstrahlung and $\text{SD}^+$ term. See the “Note on $\pi^\pm \rightarrow \ell^\pm\nu\gamma$ and $K^\pm \rightarrow \ell^\pm\nu\gamma$ Form Factors” in the $\pi^\pm$ section of the Particle Data Listings above.					
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN		
<b>&lt;2.7</b>	90	AKIBA	85	SPEC	

$\Gamma(\mu^+\nu_\mu\gamma(\text{SD}^- + \text{SD}^-\text{INT}))/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
Sum of structure-dependent part with $-\gamma$ helicity ( $\text{SD}^-$ term) and interference term between internal Bremsstrahlung and $\text{SD}^-$ term. See the “Note on $\pi^\pm \rightarrow \ell^\pm\nu\gamma$ and $K^\pm \rightarrow \ell^\pm\nu\gamma$ Form Factors” in the $\pi^\pm$ section of the Particle Data Listings above.					
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN		
<b>&lt;2.6</b>	90	<sup>1</sup> AKIBA	85	SPEC	
<sup>1</sup> Assumes $\mu$ -e universality and uses constraints from $K \rightarrow e\nu\gamma$ .					

$\Gamma(e^+\nu_e\gamma)/\Gamma(\mu^+\nu_\mu)$					$\Gamma_{16}/\Gamma_2$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1.483\pm0.066\pm0.013</math></b>	1.4K	<sup>1</sup> AMBROSINO	09E	KLOE	± $E_\gamma$ in 10–250 MeV, $p_e > 200$ MeV/c
<sup>1</sup> AMBROSINO 09E measured the differential width $dR_\gamma/dE_\gamma = (1/\Gamma(K \rightarrow \mu\nu))(d\Gamma(K \rightarrow e\nu\gamma)/dE_\gamma)$ . Result obtained by integrating the differential width over $E_\gamma$ from 10 to 250 MeV.					

$\Gamma(\pi^0e^+\nu_e\gamma)/\Gamma(\pi^0e^+\nu_e)$					$\Gamma_{17}/\Gamma_3$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.505\pm0.032</math> OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.				
0.47 ±0.02 ±0.03	4476	<sup>1</sup> AKIMENKO	07	ISTR	— $E_\gamma > 10$ MeV, $0.6 < \cos(\theta_{e\gamma}) < 0.9$
0.46 ±0.08	82	<sup>2</sup> BARMIN	91	XEBC	$E_\gamma > 10$ MeV, $0.6 < \cos(\theta_{e\gamma}) < 0.9$
0.56 ±0.04	192	<sup>3</sup> BOLOTOV	86B	CALO	— $E_\gamma > 10$ MeV
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
1.81 ±0.03 ±0.07	4476	<sup>1</sup> AKIMENKO	07	ISTR	— $E_\gamma > 10$ MeV, $\theta_{e\gamma} > 10^\circ$
0.63 ±0.02 ±0.03	4476	<sup>1</sup> AKIMENKO	07	ISTR	— $E_\gamma > 30$ MeV, $\theta_{e\gamma} > 20^\circ$
1.51 ±0.25	82	<sup>2</sup> BARMIN	91	XEBC	$E_\gamma > 10$ MeV, $\cos(\theta_{e\gamma}) < 0.98$
0.48 ±0.20	16	<sup>4</sup> LJUNG	73	HLBC	+ $E_\gamma > 30$ MeV
0.22 $\begin{smallmatrix} +0.15 \\ -0.10 \end{smallmatrix}$		<sup>4</sup> LJUNG	73	HLBC	+ $E_\gamma > 30$ MeV
0.76 ±0.28	13	<sup>5</sup> ROMANO	71	HLBC	$E_\gamma > 10$ MeV
0.53 ±0.22		<sup>5</sup> ROMANO	71	HLBC	+ $E_\gamma > 30$ MeV
1.2 ±0.8		BELLOTTI	67	HLBC	$E_\gamma > 30$ MeV

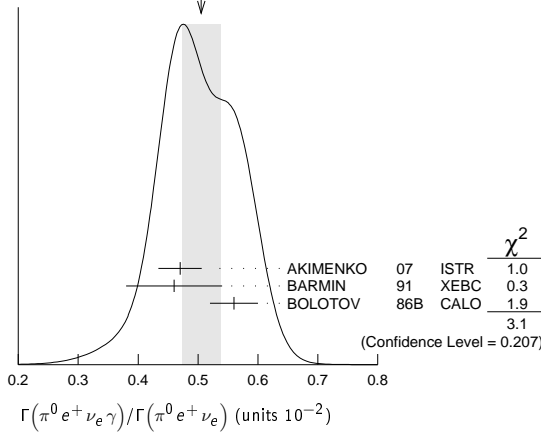
- <sup>1</sup>AKIMENKO 07 provides values for three kinematic regions. For averaging, we use value with  $E_\gamma > 10$  MeV and  $0.6 < \cos(\theta_{e\gamma}) < 0.9$ .

<sup>2</sup>BARMIN 91 quotes branching ratio  $\Gamma(K \rightarrow e\pi^0\nu\gamma)/\Gamma_{\text{all}}$ . The measured normalization is  $[\Gamma(K \rightarrow e\pi^0\nu) + \Gamma(K \rightarrow \pi^+\pi^+\pi^-)]$ . For comparison with other experiments we used  $\Gamma(K \rightarrow e\pi^0\nu)/\Gamma_{\text{all}} = 0.0482$  to calculate the values quoted here.

<sup>3</sup> $\cos(\theta_{e\gamma})$  between 0.6 and 0.9.

<sup>4</sup>First LJUNG 73 value is for  $\cos(\theta_{e\gamma}) < 0.9$ , second value is for  $\cos(\theta_{e\gamma})$  between 0.6 and 0.9 for comparison with ROMANO 71.

<sup>5</sup>Both ROMANO 71 values are for  $\cos(\theta_{e\gamma})$  between 0.6 and 0.9. Second value is for comparison with second LJUNG 73 value. We use lowest  $E_\gamma$  cut for Summary Table value. See ROMANO 71 for  $E_\gamma$  dependence.

WEIGHTED AVERAGE  
0.505±0.032 (Error scaled by 1.3) $\Gamma(\pi^0 e^+ \nu_e \gamma(\text{SD}))/\Gamma_{\text{total}}$  $\Gamma_{18}/\Gamma$ 

Structure-dependent part.

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;5.3</b>	90	BOLOTOV	86B	CALO	—

 $\Gamma(\pi^0 \mu^+ \nu_\mu \gamma)/\Gamma_{\text{total}}$  $\Gamma_{19}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.25 ± 0.25 OUR AVERAGE</b>						
1.10 ± 0.32 ± 0.05	23		<sup>1</sup> ADLER	10	B787	30 < $E_\gamma$ < 60 MeV
1.46 ± 0.22 ± 0.32	153		<sup>2</sup> TCHIKILEV	07	ISTR	— 30 < $E_\gamma$ < 60 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •						
2.4 ± 0.5 ± 0.6	125		SHIMIZU	06	K470	+ $E_\gamma > 30$ MeV; $\theta_{\mu\gamma} > 20^\circ$
<6.1	90	0	LJUNG	73	HLBC	+ $E(\gamma) > 30$ MeV

<sup>1</sup> Value obtained from  $B(K^+ \rightarrow \pi^0 \mu^+ \nu_\mu \gamma) = (2.51 \pm 0.74 \pm 0.12) \times 10^{-5}$  obtained in the kinematic region  $E_\gamma > 20$  MeV, and then theoretical  $K_{\mu 3\gamma}$  spectrum has been used. Also  $B(K^+ \rightarrow \pi^0 \mu^+ \nu_\mu \gamma) = (1.58 \pm 0.46 \pm 0.08) \times 10^{-5}$ , for  $E_\gamma > 30$  MeV and  $\theta_{\mu\gamma} > 20^\circ$ , was determined.

<sup>2</sup> Obtained from measuring  $B(K_{\mu 3\gamma})/B(K_{\mu 3})$  and using PDG 02 value  $B(K_{\mu 3}) = 3.27\%$ .  $B(K_{\mu 3\gamma}) = (8.82 \pm 0.94 \pm 0.86) \times 10^{-5}$  is obtained for 5 MeV <  $E_\gamma$  < 30 MeV.

 $\Gamma(\pi^0 \pi^0 e^+ \nu_e \gamma)/\Gamma_{\text{total}}$  $\Gamma_{20}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;5</b>	90	0	BARMIN	92	XEBC	+ $E_\gamma > 10$ MeV

## Hadronic modes with photons

 $\Gamma(\pi^+ \pi^0 \gamma(\text{INT}))/\Gamma_{\text{total}}$  $\Gamma_{21}/\Gamma$ 

The  $K^+ \rightarrow \pi^+ \pi^0 \gamma$  differential decay rate can be described in terms of  $T_{\pi^+}$ , the charged pion kinetic energy, and  $W^2 = (P_{K^+} \cdot P_\gamma) (P_{\pi^+} \cdot P_\gamma) / (m_K m_{\pi^+})^2$ ; then we can write  $d^2\Gamma(K^+ \rightarrow \pi^+ \pi^0 \gamma) / (dT_{\pi^+} dW^2) = d^2\Gamma(K^+ \rightarrow \pi^+ \pi^0 \gamma)_{IB} / (dT_{\pi^+} dW^2) [1 + 2 \cos(\pm\phi + \delta_1^+ - \delta_0^0) m_\pi^2 m_K^2 W^2 X_E + m_\pi^4 m_K^4 (X_E^2 + X_M^2) W^4]$ . The IB differential and total branching ratios are expressed in terms of the non-radiative experimental width  $\Gamma(K^+ \rightarrow \pi^+ \pi^0)$  by Low's theorem. Using PDG 10  $B(K^+ \rightarrow \pi^+ \pi^0) = 0.2066 \pm 0.0008$ , one obtains respectively  $B(K^+ \rightarrow \pi^+ \pi^0 \gamma)_{IB}$  (55 <  $T_{\pi^+}$  < 90 MeV) =  $2.55 \times 10^{-4}$  and  $B(K^+ \rightarrow \pi^+ \pi^0 \gamma)_{IB}$  (0 <  $T_{\pi^+}$  < 80 MeV) =  $1.80 \times 10^{-4}$ . Fitting respectively the piece proportional to  $W^2$  and the piece proportional to  $W^4$ , the interference contribution (INT), proportional to  $X_E$ , and the direct contribution (DE) proportional to  $X_E^2 + X_M^2$  are extracted.

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-4.24 ± 0.63 ± 0.70</b>	600k	<sup>1</sup> BATLEY	10A	NA48	± $T_{\pi^+}$ 0–80 MeV

<sup>1</sup> The cut on the photon energy implies  $W^2 > 0.2$ . BATLEY 10A obtains the INT and DE fractional branchings with respect to IB from a simultaneous kinematical fit of INT and DE and then we use the PDG 10 value for  $B(K^+ \rightarrow \pi^+ \pi^0) = 20.66 \pm 0.08$  to determine the IB. The INT and DE correlation coefficients  $-0.93$ . Assuming a constant electric amplitude,  $X_E$ , this INT value implies  $X_E = -24 \pm 6 \text{ GeV}^{-4}$ .

 $\Gamma(\pi^+ \pi^0 \gamma(\text{DE}))/\Gamma_{\text{total}}$  $\Gamma_{22}/\Gamma$ 

Direct emission (DE) part of  $\Gamma(\pi^+ \pi^0 \gamma)/\Gamma_{\text{total}}$ , assuming that interference (INT) component is zero.

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>5.99 ± 0.27 ± 0.25</b>	600k	<sup>1</sup> BATLEY	10A	NA48	± $T_{\pi^+}$ 0–80 MeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.8 ± 0.8 ± 0.7	10k	ALIEV	06	K470	+	$T_{\pi^+}$ 55–90 MeV
3.7 ± 3.9 ± 1.0	930	UVAROV	06	ISTR	—	$T_{\pi^+}$ 55–90 MeV
3.2 ± 1.3 ± 1.0	4k	ALIEV	03	K470	+	$T_{\pi^+}$ 55–90 MeV
6.1 ± 2.5 ± 1.9	4k	ALIEV	03	K470	+	$T_{\pi^+}$ full range
4.7 ± 0.8 ± 0.3	20k	<sup>2</sup> ADLER	00c	B787	+	$T_{\pi^+}$ 55–90 MeV
20.5 ± 4.6 ± 3.9		BOLOTOV	87	WIRE	—	$T_{\pi^+}$ 55–90 MeV
15.6 ± 3.5 ± 5.0		ABRAMS	72	ASPK	±	$T_{\pi^\pm}$ 55–90 MeV

<sup>1</sup> The cut on the photon energy implies  $W^2 > 0.2$ . BATLEY 10A obtains the INT and DE fractional branchings with respect to IB from a simultaneous kinematical fit of INT and DE and then we use the PDG 10 value for  $B(K^+ \rightarrow \pi^+ \pi^0) = 20.66 \pm 0.08$  to determine the IB. The INT and DE correlation coefficients  $-0.93$ . Assuming constant electric and magnetic amplitudes,  $X_E$  and  $X_M$ , these INT and DE values imply  $X_E = -24 \pm 6 \text{ GeV}^{-4}$  and  $X_M = -254 \pm 9 \text{ GeV}^{-4}$ .

<sup>2</sup> ADLER 00c measures the INT component to be  $(-0.4 \pm 1.6)\%$  of the inner bremsstrahlung (IB) component.

 $\Gamma(\pi^+ \pi^0 \pi^0 \gamma)/\Gamma(\pi^+ \pi^0 \pi^0)$  $\Gamma_{23}/\Gamma_{10}$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	CHG	COMMENT
<b>4.3 ± 3.2</b>				
-1.7	BOLOTOV	85	SPEC	— $E(\gamma) > 10$ MeV

 $\Gamma(\pi^+ \pi^+ \pi^- \gamma)/\Gamma_{\text{total}}$  $\Gamma_{24}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.04 ± 0.31 OUR AVERAGE</b>					
1.10 ± 0.48	7	BARMIN	89	XEBC	$E(\gamma) > 5$ MeV
1.0 ± 0.4		STAMER	65	EMUL	+ $E(\gamma) > 11$ MeV

 $\Gamma(\pi^+ \gamma \gamma)/\Gamma_{\text{total}}$  $\Gamma_{25}/\Gamma$ 

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>10.1 ± 0.6 OUR AVERAGE</b>						
10.03 ± 0.51 ± 0.24	215		<sup>1</sup> LAZZERONI	14	NA62	±
11 ± 3 ± 1	31		<sup>2</sup> KITCHING	97	B787	+
• • • We do not use the following data for averages, fits, limits, etc. • • •						
9.10 ± 0.72 ± 0.22	149		<sup>3</sup> BATLEY	14	NA48	±
< 0.083	90		<sup>4</sup> ARTAMONOV	05	B949	+ $P_\pi > 213$ MeV/c
< 10	90	0	ATIYA	90b	B787	+ $T_\pi$ 117–127 MeV
< 84	90	0	ASANO	82	CNTR	+ $T_\pi$ 117–127 MeV
-420 ± 520	0		ABRAMS	77	SPEC	+ $T_\pi$ < 92 MeV
< 350	90	0	LJUNG	73	HLBC	+ 6–102, 114–127 MeV
< 500	90	0	KLEMS	71	OSPK	+ $T_\pi$ < 117 MeV
-100 ± 600			CHEN	68	OSPK	+ $T_\pi$ 60–90 MeV

<sup>1</sup> LAZZERONI 14 combines NA62 and NA48/2 results. The result for the full kinematic range is extrapolated from the model-independent branching fraction  $(9.65 \pm 0.61 \pm 0.14) \times 10^{-7}$  for  $(m_{\gamma\gamma}/m_K)^2 > 0.2$ . The measured ChPT parameter  $\hat{c} = 1.86 \pm 0.25$ .

<sup>2</sup> KITCHING 97 is extrapolated from their model-independent branching fraction  $(6.0 \pm 1.5 \pm 0.7) \times 10^{-7}$  for 100 MeV/c <  $P_\pi$  < 180 MeV/c using Chiral Perturbation Theory.

<sup>3</sup> BATLEY 14 uses data collected in 2003 and 2004. Branching ratio is obtained by determining the parameter  $\hat{c} = 1.41 \pm 0.38 \pm 0.11$  and integrating the  $\mathcal{O}(p^6)$  chiral spectrum. A model independent value for the branching ratio is also obtained  $(8.77 \pm 0.87 \pm 0.17) \times 10^{-7}$  for kinematic range  $(m_{\gamma\gamma}/m_K)^2 > 0.2$ .

<sup>4</sup> ARTAMONOV 05 limit assumes ChPT with  $\hat{c} = 1.8$  with unitarity corrections. With  $\hat{c} = 1.6$  and no unitarity corrections they obtain  $< 2.3 \times 10^{-8}$  at 90% CL. This partial branching ratio is predicted to be  $6.10 \times 10^{-9}$  and  $0.49 \times 10^{-9}$  for the cases with and without unitarity correction.

 $\Gamma(\pi^+ 3\gamma)/\Gamma_{\text{total}}$  $\Gamma_{26}/\Gamma$ 

Values given here assume a phase space pion energy spectrum.

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;1.0</b>	90	ASANO	82	CNTR	+ $T(\pi)$ 117–127 MeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<3.0	90	KLEMS	71	OSPK	+ $T(\pi)$ > 117 MeV
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 $\Gamma(\pi^+ e^+ e^- \gamma)/\Gamma_{\text{total}}$  $\Gamma_{27}/\Gamma$ 

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.19 ± 0.12 ± 0.04</b>	113	<sup>1</sup> BATLEY	08	NA48 $m_{ee\gamma} > 260$ MeV

<sup>1</sup> BATLEY 08 also reports the Chiral Perturbation Theory parameter  $\hat{c} = 0.9 \pm 0.45$  obtained using the shape of the  $e^+ e^- \gamma$  invariant mass spectrum. By extrapolating the theoretical amplitude to  $m_{ee\gamma} < 260$  MeV, it obtains the inclusive  $B(K^+ \rightarrow \pi^+ e^+ e^- \gamma) = (1.29 \pm 0.13 \pm 0.03) \times 10^{-8}$ , where the first error is the combined statistical and systematic errors and the second error is from the uncertainty in  $\hat{c}$ .

Leptonic modes with  $\ell\bar{\ell}$  pairs $\Gamma(e^+ \nu_e \nu_\ell)/\Gamma(e^+ \nu_e)$  $\Gamma_{28}/\Gamma_1$ 

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;3.8</b>	90	0	HEINTZE	79	SPEC +

 $\Gamma(\mu^+ \nu_\mu \nu_\ell)/\Gamma_{\text{total}}$  $\Gamma_{29}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;6.0</b>	90	0	<sup>1</sup> PANG	73	CNTR +

<sup>1</sup> PANG 73 assumes  $\mu$  spectrum from  $\nu$ - $\nu$  interaction of BARDIN 70.

# Meson Particle Listings

$K^\pm$

$\Gamma(e^+ \nu_e e^+ e^-)/\Gamma_{\text{total}}$						$\Gamma_{30}/\Gamma$	
VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
<b><math>2.48 \pm 0.14 \pm 0.14</math></b>		410	POBLAGUEV 02	B865	+	$m_{ee} > 150$ MeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
20 $\pm 20$		4	DIAMANT-...	76	SPEC	$m_{e^+e^-} > 140$ MeV	

$\Gamma(\mu^+ \nu_\mu e^+ e^-)/\Gamma_{\text{total}}$						$\Gamma_{31}/\Gamma$	
VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
<b><math>7.06 \pm 0.16 \pm 0.26</math></b>		2.7k	POBLAGUEV 02	B865	+	$m_{ee} > 145$ MeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
100 $\pm 30$		14	DIAMANT-...	76	SPEC	$m_{e^+e^-} > 140$ MeV	

$\Gamma(e^+ \nu_e \mu^+ \mu^-)/\Gamma_{\text{total}}$						$\Gamma_{32}/\Gamma$	
VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
<b><math>1.72 \pm 0.45</math></b>			MA 06	B865			
• • • We do not use the following data for averages, fits, limits, etc. • • •							
<50		90	ADLER 98	B787			

$\Gamma(\mu^+ \nu_\mu \mu^+ \mu^-)/\Gamma_{\text{total}}$						$\Gamma_{33}/\Gamma$	
VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
<b>&lt;4.1</b>		90	ATIYA 89	B787	+		

Lepton Family number (LF), Lepton number (L),  $\Delta S = \Delta Q$  (SQ) violating modes, or  $\Delta S = 1$  weak neutral current (SI) modes

$\Gamma(\pi^+ \pi^+ e^- \bar{\nu}_e)/\Gamma_{\text{total}}$						$\Gamma_{34}/\Gamma$	
Test of $\Delta S = \Delta Q$ rule.							
VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
< 9.0	95	0	SCHWEINB...	71	HLBC	+	
< 6.9	95	0	ELY	69	HLBC	+	
<20.	95		BIRGE	65	FBC	+	

$\Gamma(\pi^+\pi^+e^-\bar{\nu}_e)/\Gamma(\pi^+\pi^-e^+\nu_e)$					$\Gamma_{34}/\Gamma_6$
Test of $\Delta S = \Delta Q$ rule.					
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	
<b>&lt; 3</b>	90	3	<sup>1</sup> BLOCH	76	SPEC
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<130.	95	0	BOURQUIN	71	ASPK
<sup>1</sup> BLOCH 76 quotes $3.6 \times 10^{-4}$ at CL = 95%, we convert.					

<sup>1</sup> BLOCH 76 quotes  $3.6 \times 10^{-4}$  at CL = 95%, we convert.

$\Gamma(\pi^+ \pi^+ \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$				$\Gamma_{35}/\Gamma$		
Test of $\Delta S = \Delta Q$ rule.						
VALUE (units $10^{-6}$ )	CL%	EVTS		DOCUMENT ID	TECN	CHG
<3.0	95	0		BIRGE 65	FBC	+

$$\Gamma(\pi^+ e^+ e^-)/\Gamma_{\text{total}}$$

$$\Gamma_{36}/\Gamma$$

Test for  $\Delta S = 1$  weak neutral current. Allowed by combined first-order weak and electromagnetic interactions.

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b><math>3.00 \pm 0.09</math> OUR AVERAGE</b>				
$3.11 \pm 0.04 \pm 0.12$	7253	<sup>1</sup> BATLEY 09	NA48	$\pm$
$2.94 \pm 0.05 \pm 0.14$	10300	<sup>2</sup> APPEL 99	SPEC	+
$2.75 \pm 0.23 \pm 0.13$	500	<sup>3</sup> ALLIEGRO 92	SPEC	+
$2.7 \pm 0.5$	41	<sup>4</sup> BLOCH 75	SPEC	+

<sup>1</sup> Value extrapolated from a measurement in the region  $z = (m_{ee}/m_K)^2 > 0.08$ . BATLEY 09 also evaluated the shape of the form factor using four different theoretical models.

<sup>2</sup> APPEL 99 establishes vector nature of this decay and determines form factor  $f(z) = f_0(1+\delta z)$ ,  $Z = M_{ee}^2/m_K^2$ ,  $\delta = 2.14 \pm 0.13 \pm 0.15$ .

<sup>3</sup> ALLIEGRO 92 assumes a vector interaction with a form factor given by  $\lambda = 0.105 \pm 0.035 \pm 0.015$  and a correlation coefficient of  $-0.82$ .

<sup>4</sup> BLOCH 75 assumes a vector interaction.

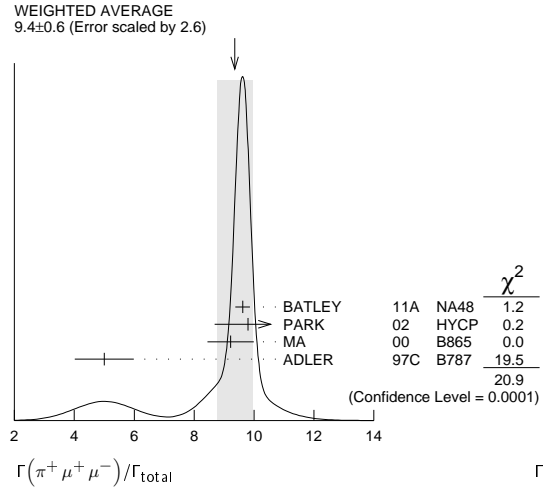
$\Gamma(\pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$						$\Gamma_{37}/\Gamma$
Test for $\Delta S = 1$ weak neutral current. Allowed by higher-order electroweak interactions.						
VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG.	COMMENT
<b><math>9.4 \pm 0.6</math></b>	<b>OUR AVERAGE</b>		Error includes scale factor of 2.6. See the ideogram below.			
$9.62 \pm 0.21 \pm 0.13$		3120	<sup>1</sup> BATLEY	11A	NA48	$\pm$ 2003-04 data
$9.8 \pm 1.0 \pm 0.5$		110	<sup>2</sup> PARK	02	HYCP	$\pm$
$9.22 \pm 0.60 \pm 0.49$		402	<sup>3</sup> MA	00	B865	+
$5.0 \pm 0.4 \pm 0.9$		207	<sup>4</sup> ADLER	97c	B787	+
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
$9.7 \pm 1.2 \pm 0.4$		65	PARK	02	HYCP	+
$10.0 \pm 1.9 \pm 0.7$		35	PARK	02	HYCP	-
<23		90	ATIYA	89	B787	+

<sup>1</sup> BATLEY 11A also studies the form factor  $f(z)$  dependence of the decay, described via single photon exchange: i) assuming a linear form factor,  $f(z) = f_0(1 + \delta z)$ ,  $z = (M_{\mu\mu}/m_K)^2$ , finding  $f_0 = 0.470 \pm 0.040$  and  $\delta = 3.11 \pm 0.57$  and ii) assuming a linear form factor including  $\pi$ - $\pi$  rescattering,  $W_{\pi\pi}$ , as in DAMBROSIO 98a, finding  $f(z) = G_F m_K^2 (a_+ + b_+ z) + W_{\pi\pi}(z)$ ,  $a_+ = -0.575 \pm 0.039$ ,  $b_+ = -0.813 \pm 0.145$ .

<sup>2</sup> PARK 02 “ $\pm$ ” result comes from combining  $K^+ \rightarrow \pi^+ \mu^+ \mu^-$  and  $K^- \rightarrow \pi^- \mu^+ \mu^-$ , assuming  $CP$  is conserved.

<sup>3</sup> MA 00 establishes vector nature of this decay and determines form factor  $f(z) = f_0(1 + \delta z)$ ,  $z = (M_{\mu\mu}/m_K)^2$ ,  $\delta = 2.45 \pm 1.30 - 0.95$ .

<sup>4</sup> ADLER 97c gives systematic error  $0.7 \times 10^{-8}$  and theoretical uncertainty  $0.6 \times 10^{-8}$ , which we combine in quadrature to obtain our second error.



$\Gamma(\pi^+\nu\bar{\nu})/\Gamma_{\text{total}}$	$\Gamma_{38}/\Gamma$
Test for $\Delta S = 1$ weak neutral current. Allowed by higher-order electroweak interactions. Branching ratio values are extrapolated from the momentum or energy regions shown in the comments assuming Standard Model phase space except for those labeled “Scalar” or “Tensor” to indicate the assumed non-Standard-Model interaction.	

• • • We do not use the following data for averages, fits, limits, etc. • • •							
0.789 $\pm 0.926 - 0.510$		3	<sup>2</sup> ARTAMONOV 08	B949	+	$140 < P_\pi < 199$ MeV	
< 2.2		90	<sup>3</sup> ADLER 04	B787	+	$211 < P_\pi < 229$ MeV	
< 2.7		90	ADLER 04	B787	+	Scalar	
< 1.8		90	ADLER 04	B787	+	Tensor	
0.147 $\pm 0.130 - 0.089$		3	<sup>4</sup> ANISIMOVSKY..04	B949	+	$211 < P_\pi < 229$ MeV	
0.157 $\pm 0.175 - 0.082$		2	ADLER 02	B787	+	$P_\pi > 211$ MeV/c	
< 4.2		90	ADLER 02c	B787	+	$140 < P_\pi < 195$ MeV	
< 4.7		90	<sup>5</sup> ADLER 02c	B787	+	Scalar	
< 2.5		90	<sup>5</sup> ADLER 02c	B787	+	Tensor	
0.15 $\pm 0.34 - 0.12$		1	ADLER 00	B787		In ADLER 02	
0.42 $\pm 0.97 - 0.35$		1	ADLER 97	B787			
< 2.4		90	ADLER 96	B787			
< 7.5		90	ATIYA 93	B787	+	$T(\pi) 115-127$ MeV	
< 5.2		90	<sup>6</sup> ATIYA 93	B787	+		
< 17		90	ATIYA 93b	B787	+	$T(\pi) 60-100$ MeV	
< 34		90	ATIYA 90	B787	+		
<140		90	ASANO 81b	CNTR	+	$T(\pi) 116-127$ MeV	

<sup>1</sup> Value obtained combining ANISIMOVSKY 04, ADLER 04, and the present ARTAMONOV 08 results.

<sup>2</sup> Observed 3 events with an estimated background of  $0.93 \pm 0.17 \pm 0.32 - 0.24$ . Signal-to-background ratio for each of these 3 events is 0.20, 0.42, and 0.47.

<sup>3</sup> Value obtained combining the previous result ADLER 02c with 1 event and the present result with 0 events to obtain an expected background  $1.22 \pm 0.24$  events and 1 event observed.

<sup>4</sup> Value obtained combining the previous E787 result ADLER 02 with 2 events and the present E949 with 1 event. The additional event has a signal-to-background ratio 0.9. Superseded by ARTAMONOV 08.

<sup>5</sup> Superseded by ADLER 04.

<sup>6</sup> Combining ATIYA 93 and ATIYA 93b results. Superseded by ADLER 96.

$\Gamma(\pi^+\pi^0\nu\bar{\nu})/\Gamma_{\text{total}}$					$\Gamma_{39}/\Gamma$
Test for $\Delta S = 1$ weak neutral current. Allowed by higher-order electroweak interactions.					
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	
<b>&lt;4.3</b>	90	<sup>1</sup>	ADLER	01	SPEC

<sup>1</sup> Search region defined by  $90 \text{ MeV}/c < P_{\pi^+} < 188 \text{ MeV}/c$  and  $135 \text{ MeV} < E_{\pi^0} < 180 \text{ MeV}$ .

$\Gamma(\mu^- \nu e^+ e^+)/\Gamma(\pi^+ \pi^- e^+ \nu_e)$							$\Gamma_{40}/\Gamma_6$
Test of lepton family number conservation.							
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG		
<0.5	90	0	<sup>1</sup> DIAMANT-...	76	SPEC	+	

<sup>1</sup> DIAMANT-BERGER 76 quotes this result times our 1975  $\pi^+ \pi^- e \nu$  BR ratio.

See key on page 601

## Meson Particle Listings

 $K^\pm$  $\Gamma(\mu^+\nu_e)/\Gamma_{\text{total}}$   $\Gamma_{41}/\Gamma$ 

Forbidden by lepton family number conservation.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.004</b>	90	0	<sup>1</sup> LYONS 81	HLBC	200 GeV $K^+$ narrow band $\nu$ beam

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.012 90 <sup>1</sup> COOPER 82 HLBC Wideband  $\nu$  beam<sup>1</sup> COOPER 82 and LYONS 81 limits on  $\nu_e$  observation are here interpreted as limits on lepton family number violation in the absence of mixing. $\Gamma(\pi^+\mu^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{42}/\Gamma$ 

Test of lepton family number conservation.

VALUE (units $10^{-10}$ )	CL%	DOCUMENT ID	TECN	CHG
<b>&lt;0.13</b>	90	<sup>1</sup> SHER 05	RVUE	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

&lt;0.21 90 SHER 05 B865 +

&lt;0.39 90 APPEL 00 B865 +

&lt;2.1 90 LEE 90 SPEC +

<sup>1</sup> This result combines SHER 05 1998 data, APPEL 00 1996 data, and data from BERGMAN 97 and PISLAK 97 theses, all from BNL-E865, with LEE 90 BNL-E777 data. $\Gamma(\pi^+\mu^-e^+)/\Gamma_{\text{total}}$   $\Gamma_{43}/\Gamma$ 

Test of lepton family number conservation.

VALUE (units $10^{-10}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt; 5.2</b>	90	0	APPEL 00B	B865	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

<70 90 <sup>1</sup> DIAMANT-... 76 SPEC +<sup>1</sup> Measurement actually applies to the sum of the  $\pi^+\mu^-e^+$  and  $\pi^-\mu^+e^+$  modes. $\Gamma(\pi^-\mu^+e^+)/\Gamma_{\text{total}}$   $\Gamma_{44}/\Gamma$ 

Test of total lepton number conservation.

VALUE (units $10^{-10}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt; 5.0</b>	90	0	APPEL 00B	B865	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

<70 90 <sup>1</sup> DIAMANT-... 76 SPEC +<sup>1</sup> Measurement actually applies to the sum of the  $\pi^+\mu^-e^+$  and  $\pi^-\mu^+e^+$  modes. $\Gamma(\pi^-e^+e^+)/\Gamma_{\text{total}}$   $\Gamma_{45}/\Gamma$ 

Test of total lepton number conservation.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;6.4 <math>\times 10^{-10}</math></b>	90	0	APPEL 00B	B865	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

<9.2  $\times 10^{-9}$  90 <sup>1</sup> DIAMANT-... 76 SPEC +<1.5  $\times 10^{-5}$  90 CHANG 68 HBC - $\Gamma(\pi^-\mu^+\mu^+)/\Gamma_{\text{total}}$   $\Gamma_{46}/\Gamma$ 

Forbidden by total lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	CHG
<b>&lt;1.1 <math>\times 10^{-9}</math></b>	90	BATLEY 11A	NA48	$\pm$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<3.0  $\times 10^{-9}$  90 APPEL 00B B865 +<1.5  $\times 10^{-4}$  90 <sup>1</sup> LITTENBERG 92 HBC<sup>1</sup> LITTENBERG 92 is from retroactive data analysis of CHANG 68 bubble chamber data. $\Gamma(\mu^+\bar{\nu}_e)/\Gamma_{\text{total}}$   $\Gamma_{47}/\Gamma$ 

Forbidden by total lepton number conservation.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.3</b>	90	<sup>1</sup> COOPER 82	HLBC	Wideband $\nu$ beam

<sup>1</sup> COOPER 82 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton number violation in the absence of mixing. $\Gamma(\pi^0e^+\bar{\nu}_e)/\Gamma_{\text{total}}$   $\Gamma_{48}/\Gamma$ 

Forbidden by total lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.003</b>	90	<sup>1</sup> COOPER 82	HLBC	Wideband $\nu$ beam

<sup>1</sup> COOPER 82 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton number violation in the absence of mixing. $\Gamma(\pi^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{49}/\Gamma$ 

Violates angular momentum conservation and gauge invariance. Current interest in this decay is as a search for non-commutative space-time effects as discussed in ARTAMONOV 05 and for exotic physics such as a vacuum expectation value of a new vector field, non-local Superstring effects, or departures from Lorentz invariance, as discussed in ADLER 02b.

VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID	TECN	CHG
<b>&lt; 2.3</b>	90	ARTAMONOV 05	B949	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

&lt; 360 90 ADLER 02b B787 +

&lt;1400 90 ASANO 82 CNTR +

<4000 90 <sup>1</sup> KLEMS 71 OSPK +<sup>1</sup> Test of model of Selleri, Nuovo Cimento **60A** 291 (1969).CPT VIOLATION TESTS IN  $K^\pm$  DECAYS

$$\Delta = (\Gamma(K^+) - \Gamma(K^-)) / (\Gamma(K^+) + \Gamma(K^-))$$

 $\Delta(K^\pm \rightarrow \mu^\pm \nu_\mu)$  RATE DIFFERENCE/SUM

VALUE (%)	DOCUMENT ID	TECN
<b>-0.27 <math>\pm</math> 0.21</b>	FORD	67 CNTR

 $\Delta(K^\pm \rightarrow \pi^\pm \pi^0)$  RATE DIFFERENCE/SUM

VALUE (%)	DOCUMENT ID	TECN
<b>0.4 <math>\pm</math> 0.6</b>	HERZO	69 OSPK

CP VIOLATION TESTS IN  $K^\pm$  DECAYS

$$\Delta = (\Gamma(K^+) - \Gamma(K^-)) / (\Gamma(K^+) + \Gamma(K^-))$$

 $\Delta(K^\pm \rightarrow \pi^\pm e^+ e^-)$  RATE DIFFERENCE/SUM

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN
<b>-2.2 <math>\pm</math> 1.5 <math>\pm</math> 0.6</b>	<sup>1</sup> BATLEY 09	NA48

<sup>1</sup> This implies an upper limit of  $2.1 \times 10^{-2}$  at 90% CL. $\Delta(K^\pm \rightarrow \pi^\pm \mu^+ \mu^-)$  RATE DIFFERENCE/SUM

VALUE	DOCUMENT ID	TECN
<b>0.010 <math>\pm</math> 0.023 OUR AVERAGE</b>		
0.011 $\pm$ 0.023	<sup>1</sup> BATLEY 11A	NA48
-0.02 $\pm$ 0.11 $\pm$ 0.04	PARK 02	HYCP

<sup>1</sup> This corresponds to the asymmetry upper limit of  $< 2.9 \times 10^{-2}$  at 90% CL. $\Delta(K^\pm \rightarrow \pi^\pm \pi^0 \gamma)$  RATE DIFFERENCE/SUM

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.0 <math>\pm</math> 1.2 OUR AVERAGE</b>					
0.0 $\pm$ 1.0 $\pm$ 0.6	1M	<sup>1</sup> BATLEY 10A	NA48		
4 $\pm$ 29	2461	SMITH 76	WIRE $\pm$		$E_\pi$ 55-90 MeV
5 $\pm$ 20	4000	ABRAMS 73B	ASPK $\pm$		$E_\pi$ 51-100 MeV

<sup>1</sup> This value implies the upper bound for this asymmetry  $1.5 \times 10^{-3}$  at 90% CL. $\Delta(K^\pm \rightarrow \pi^\pm \pi^+ \pi^-)$  RATE DIFFERENCE/SUM

VALUE (%)	EVTS	DOCUMENT ID	TECN	CHG
<b>0.04 <math>\pm</math> 0.06</b>		<sup>1</sup> FORD 70	ASP K	
-0.01 $\pm$ 0.08		<sup>2</sup> SMITH 73	ASP K $\pm$	
0.05 $\pm$ 0.07	3.2M	<sup>1</sup> FORD 70	ASP K	
-0.25 $\pm$ 0.45		FLETCHER 67	OSPK	
-0.02 $\pm$ 0.11		<sup>1</sup> FORD 67	CNTR	

<sup>1</sup> First FORD 70 value is second FORD 70 combined with FORD 67.<sup>2</sup> SMITH 73 value of  $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$  rate difference is derived from SMITH 73 value of  $K^\pm \rightarrow \pi^\pm 2\pi^0$  rate difference. $\Delta(K^\pm \rightarrow \pi^\pm \pi^0 \pi^0)$  RATE DIFFERENCE/SUM

VALUE (%)	EVTS	DOCUMENT ID	TECN	CHG
<b>-0.02 <math>\pm</math> 0.28 OUR AVERAGE</b>				
0.04 $\pm$ 0.29		SMITH 73	ASPK $\pm$	
-0.6 $\pm$ 0.9	1802	HERZO 69	OSPK	

T VIOLATION TESTS IN  $K^+$  AND  $K^-$  DECAYS $P_T$  in  $K^+ \rightarrow \pi^0 \mu^+ \nu_\mu$ 

T-violating muon polarization. Sensitive to new sources of CP violation beyond the Standard Model.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>-1.7 <math>\pm</math> 2.3 <math>\pm</math> 1.1</b>		<sup>1</sup> ABE 04F	K246	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

-4.2  $\pm$  4.9  $\pm$  0.9 3.9M ABE 99s K246 +<sup>1</sup> Includes three sets of data: 96-97 (ABE 99s), 98, and 99-00 totaling about three times the ABE 99s data sample. Corresponds to  $P_T < 5.0 \times 10^{-3}$  at 90% CL. $P_T$  in  $K^+ \rightarrow \mu^+ \nu_\mu \gamma$ 

T-violating muon polarization. Sensitive to new sources of CP violation beyond the Standard Model.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>-0.64 <math>\pm</math> 1.85 <math>\pm</math> 0.10</b>	114k	<sup>1</sup> ANISIMOVSK...03	K246	+

<sup>1</sup> Muons stopped and polarization measured from decay to positrons. $\text{Im}(\xi)$  in  $K^+ \rightarrow \pi^0 \mu^+ \nu_\mu$  DECAY (from transverse  $\mu$  pol.)

Test of T reversal invariance.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.006 <math>\pm</math> 0.008 OUR AVERAGE</b>					
-0.0053 $\pm$ 0.0071 $\pm$ 0.0036		<sup>1</sup> ABE 04F	K246	+	
-0.016 $\pm$ 0.025	20M	CAMPBELL 81	CNTR	+	Pol.
• • • We do not use the following data for averages, fits, limits, etc. • • •					
-0.013 $\pm$ 0.016 $\pm$ 0.003	3.9M	ABE 99s	CNTR	+	$p_T$ $K^+$ at rest

<sup>1</sup> Includes three sets of data: 96-97 (ABE 99s), 98, and 99-00 totaling about three times the ABE 99s data sample. Corresponds to  $\text{Im}(\xi) < 0.016$  at 90% CL.

# Meson Particle Listings

$K^\pm$

## DALITZ PLOT PARAMETERS FOR $K \rightarrow 3\pi$ DECAYS

Revised 1999 by T.G. Trippe (LBNL).

The Dalitz plot distribution for  $K^\pm \rightarrow \pi^\pm \pi^\pm \pi^\mp$ ,  $K^\pm \rightarrow \pi^0 \pi^0 \pi^\pm$ , and  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$  can be parameterized by a series expansion such as that introduced by Weinberg [1]. We use the form

$$\begin{aligned} |M|^2 &\propto 1 + g \frac{(s_3 - s_0)}{m_{\pi^+}^2} + h \left[ \frac{s_3 - s_0}{m_{\pi^+}^2} \right]^2 \\ &+ j \frac{(s_2 - s_1)}{m_{\pi^+}^2} + k \left[ \frac{s_2 - s_1}{m_{\pi^+}^2} \right]^2 \\ &+ f \frac{(s_2 - s_1)(s_3 - s_0)}{m_{\pi^+}^2 m_{\pi^+}^2} + \dots, \end{aligned} \quad (1)$$

where  $m_{\pi^+}^2$  has been introduced to make the coefficients  $g$ ,  $h$ ,  $j$ , and  $k$  dimensionless, and

$$\begin{aligned} s_i &= (P_K - P_i)^2 = (m_K - m_i)^2 - 2m_K T_i, \quad i = 1, 2, 3, \\ s_0 &= \frac{1}{3} \sum_i s_i = \frac{1}{3} (m_K^2 + m_1^2 + m_2^2 + m_3^2). \end{aligned}$$

Here the  $P_i$  are four-vectors,  $m_i$  and  $T_i$  are the mass and kinetic energy of the  $i^{\text{th}}$  pion, and the index 3 is used for the odd pion.

The coefficient  $g$  is a measure of the slope in the variable  $s_3$  (or  $T_3$ ) of the Dalitz plot, while  $h$  and  $k$  measure the quadratic dependence on  $s_3$  and  $(s_2 - s_1)$ , respectively. The coefficient  $j$  is related to the asymmetry of the plot and must be zero if  $CP$  invariance holds. Note also that if  $CP$  is good,  $g$ ,  $h$ , and  $k$  must be the same for  $K^+ \rightarrow \pi^+ \pi^+ \pi^-$  as for  $K^- \rightarrow \pi^- \pi^- \pi^+$ .

Since different experiments use different forms for  $|M|^2$ , in order to compare the experiments we have converted to  $g$ ,  $h$ ,  $j$ , and  $k$  whatever coefficients have been measured. Where such conversions have been done, the measured coefficient  $a_y$ ,  $a_t$ ,  $a_u$ , or  $a_v$  is given in the comment at the right. For definitions of these coefficients, details of this conversion, and discussion of the data, see the April 1982 version of this note [2].

## References

1. S. Weinberg, Phys. Rev. Lett. **4**, 87 (1960).
2. Particle Data Group, Phys. Lett. **111B**, 69 (1982).

## ENERGY DEPENDENCE OF $K^\pm$ DALITZ PLOT

$$\begin{aligned} |\text{matrix element}|^2 &= 1 + gu + hu^2 + kv^2 \\ \text{where } u &= (s_3 - s_0) / m_\pi^2 \text{ and } v = (s_2 - s_1) / m_\pi^2 \end{aligned}$$

## LINEAR COEFFICIENT $g$ FOR $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$

Some experiments use Dalitz variables  $x$  and  $y$ . In the comments we give  $a_y$  = coefficient of  $y$  term. See note above on "Dalitz Plot Parameters for  $K \rightarrow 3\pi$  Decays." For discussion of the conversion of  $a_y$  to  $g$ , see the earlier version of the same note in the Review published in Physics Letters **111B** 70 (1982).

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.21134 ± 0.00017</b>	471M	<sup>1</sup> BATLEY	07B	NA48	±
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
-0.2221 ± 0.0065	225k	DEVAUX	77	SPEC	+ $a_y = .2814 \pm .0082$
-0.199 ± 0.008	81k	LUCAS	73	HBC	- $a_y = 0.252 \pm 0.011$
-0.2157 ± 0.0028	750k	FORD	72	ASPK	+ $a_y = .2734 \pm .0035$
-0.2186 ± 0.0028	750k	FORD	72	ASPK	- $a_y = .2770 \pm .0035$
-0.200 ± 0.009	39819	<sup>3</sup> HOFFMASTER	72	HLBC	+
-0.196 ± 0.012	17898	<sup>4</sup> GRAUMAN	70	HLBC	+ $a_y = 0.228 \pm 0.030$

-0.193 ± 0.010	50919	MAST	69	HBC	- $a_y = 0.244 \pm 0.013$
-0.218 ± 0.016	9994	<sup>5</sup> BUTLER	68	HBC	+ $a_y = 0.277 \pm 0.020$
-0.190 ± 0.023	5778	<sup>5,6</sup> MOSCOSO	68	HBC	- $a_y = 0.242 \pm 0.029$
-0.22 ± 0.024	5428	<sup>5,6</sup> ZINCHENKO	67	HBC	+ $a_y = 0.28 \pm 0.03$
-0.220 ± 0.035	1347	<sup>7</sup> FERRO-LUZZI	61	HBC	- $a_y = 0.28 \pm 0.045$

<sup>1</sup> Final state strong interaction and radiative corrections not included in the fit.

<sup>2</sup> Quadratic dependence is required by  $K_L^0$  experiments.

<sup>3</sup> HOFFMASTER 72 includes GRAUMAN 70 data.

<sup>4</sup> Emulsion data added — all events included by HOFFMASTER 72.

<sup>5</sup> Experiments with large errors not included in average.

<sup>6</sup> Also includes DBC events.

<sup>7</sup> No radiative corrections included.

## QUADRATIC COEFFICIENT $h$ FOR $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>1.848 ± 0.040</b>	471 M	<sup>1</sup> BATLEY	07B	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.06 ± 1.43	225 k	DEVAUX	77	SPEC +
1.87 ± 0.62	750 k	FORD	72	ASPK +
1.25 ± 0.62	750 k	FORD	72	ASPK -
-0.9 ± 1.4	39819	HOFFMASTER	72	HLBC +
-0.1 ± 1.2	50919	MAST	69	HBC -

<sup>1</sup> Final state strong interaction and radiative corrections not included in the fit.

## QUADRATIC COEFFICIENT $k$ FOR $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>-4.63 ± 0.14</b>	471M	<sup>1</sup> BATLEY	07B	NA48 ±
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
-20.5 ± 3.9	225k	DEVAUX	77	SPEC +
-7.5 ± 1.9	750k	FORD	72	ASPK +
-8.3 ± 1.9	750k	FORD	72	ASPK -
-10.5 ± 4.5	39819	HOFFMASTER	72	HLBC +
-14 ± 12	50919	MAST	69	HBC -

<sup>1</sup> Final state strong interaction and radiative corrections not included in the fit.

## $(g_+ - g_-) / (g_+ + g_-)$ FOR $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$

This is a  $CP$  violating asymmetry between linear coefficients  $g_\pm$  for  $K^+ \rightarrow \pi^+ \pi^+ \pi^-$  decay and  $g_-$  for  $K^- \rightarrow \pi^- \pi^+ \pi^-$  decay.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN
<b>-1.5 ± 1.5 ± 1.6</b>	3.1G	<sup>1</sup> BATLEY	07E NA48
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
1.7 ± 2.1 ± 2.0	1.7G	<sup>2</sup> BATLEY	06 NA48
-70.0 ± 53	3.2M	FORD	70 ASPK

<sup>1</sup> BATLEY 07E includes data from BATLEY 06. Uses quadratic parametrization and value  $g_+ + g_- = 2g$  from BATLEY 07B. This measurement neglects any possible charge asymmetries in higher order slope parameters  $h$  or  $k$ .

<sup>2</sup> This measurement neglects any possible charge asymmetries in higher order slope parameters  $h$  or  $k$ .

## LINEAR COEFFICIENT $g$ FOR $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

Unless otherwise stated, all experiments include terms quadratic in  $(s_3 - s_0) / m_\pi^2$ . See note above on "Dalitz Plot Parameters for  $K \rightarrow 3\pi$  Decays."

See BATUSOV 98 for a discussion of the discrepancy between their result and others, especially BOLOTOV 86. At this time we have no way to resolve the discrepancy so we depend on the large scale factor as a warning.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.626 ± 0.007 OUR AVERAGE</b>					
0.6259 ± 0.0043 ± 0.0093	493k	AKOPDZHAN.	05B	TNF	±
0.627 ± 0.004 ± 0.010	252k	<sup>1,2</sup> AJINENKO	03B	ISTR	-
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
0.736 ± 0.014 ± 0.012	33k	BATUSOV	98	SPEC	+
0.582 ± 0.021	43k	BOLOTOV	86	CALO	-
0.670 ± 0.054	3263	BRAUN	76B	HLBC	+
0.630 ± 0.038	5635	SHEAFF	75	HLBC	+
0.510 ± 0.060	27k	SMITH	75	WIRE	+
0.67 ± 0.06	1365	AUBERT	72	HLBC	+
0.544 ± 0.048	4048	DAVISON	69	HLBC	+ Also emulsion

<sup>1</sup> Measured using in-flight decays of the 25 GeV negative secondary beam.

<sup>2</sup> They form new world averages  $g_- = (0.617 \pm 0.018)$  and  $g_+ = (0.684 \pm 0.033)$  which give  $\Delta g_{\pi^0} = 0.051 \pm 0.028$ .

## QUADRATIC COEFFICIENT $h$ FOR $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.052 ± 0.008 OUR AVERAGE</b>					
0.0551 ± 0.0044 ± 0.0086	493k	AKOPDZHAN.	05B	TNF	±
0.046 ± 0.004 ± 0.012	252k	<sup>1</sup> AJINENKO	03B	ISTR	-
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
0.128 ± 0.015 ± 0.024	33k	BATUSOV	98	SPEC	+
0.037 ± 0.024	43k	BOLOTOV	86	CALO	-
0.152 ± 0.082	3263	BRAUN	76B	HLBC	+
0.041 ± 0.030	5635	SHEAFF	75	HLBC	+
0.009 ± 0.040	27k	SMITH	75	WIRE	+
-0.01 ± 0.08	1365	AUBERT	72	HLBC	+
0.026 ± 0.050	4048	DAVISON	69	HLBC	+ Also emulsion

<sup>1</sup> Measured using in-flight decays of the 25 GeV negative secondary beam.



**QUADRATIC COEFFICIENT  $k$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b><math>0.0054 \pm 0.0035</math> OUR AVERAGE</b>		Error includes scale factor of 2.5.		
$0.0082 \pm 0.0011 \pm 0.0014$	493k	AKOPDZHAN..05b	TNF	$\pm$
$0.001 \pm 0.001 \pm 0.002$	252k	<sup>1</sup> AJINENKO	03b ISTR	$-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.0197 \pm 0.0045 \pm 0.0029$	33k	BATUSOV	98 SPEC	$+$

<sup>1</sup> Measured using in-flight decays of the 25 GeV negative secondary beam.

 **$(g_+ - g_-) / (g_+ + g_-)$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

A nonzero value for this quantity indicates  $CP$  violation.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b><math>1.8 \pm 1.8</math> OUR AVERAGE</b>				
$1.8 \pm 1.7 \pm 0.6$	91.3M	<sup>1</sup> BATLEY	07e NA48	
$2 \pm 18 \pm 5$	619k	<sup>2</sup> AKOPDZHAN..05	TNF	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.8 \pm 2.2 \pm 1.3$	47M	<sup>3</sup> BATLEY	06a NA48	

<sup>1</sup> BATLEY 07e includes data from BATLEY 06a. Uses quadratic parametrization and PDG 06 value  $g = 0.626 \pm 0.007$  to obtain  $g_+ - g_- = (2.2 \pm 2.1 \pm 0.7) \times 10^{-4}$ . Neglects any possible charge asymmetries in higher order slope parameters  $h$  or  $k$ .

<sup>2</sup> Asymmetry obtained assuming that  $g_+ + g_- = 2 \times 0.652$  (PDG 02) and that asymmetries in  $h$  and  $k$  are zero.

<sup>3</sup> Linear and quadratic slopes from PDG 04 are used. Any possible charge asymmetries in higher order slope parameters  $h$  or  $k$  are neglected.

**ALTERNATIVE PARAMETRIZATIONS OF  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$  DALITZ PLOT**

The following functional form for the matrix element suggested by  $\pi\pi$  rescattering in  $K^+ \rightarrow \pi^+ \pi^+ \pi^- \pi^0 \rightarrow \pi^+ \pi^0 \pi^0$  is used for this fit (CABIBBO 04a, CABIBBO 05): Matrix element =  $M_0 + M_1$  where  $M_0 = 1 + (1/2)g_0 u + (1/2)h' u^2 + (1/2)k_0 v^2$  with  $u = (s_3 - s_0)/(m_{\pi^+})^2$ ,  $v = (s_2 - s_1)/(m_{\pi^+})^2$  and where  $M_1$  takes into account the non-analytic piece due to  $\pi\pi$  rescattering amplitudes  $a_0$  and  $a_2$ ; The parameters  $g_0$  and  $h'$  are related to the parameters  $g$  and  $h$  of the matrix element squared given in the previous section by the approximations  $g_0 \sim g^{PDG}$  and  $h' \sim h^{PDG} - (g/2)^2$  and  $k_0 \sim k^{PDG}$ .

In addition, we also consider the effective field theory framework of COLANGELO 06a and BISSEGGER 09 to extract  $g_{BB}$  and  $h'_{BB}$ .

**LINEAR COEFFICIENT  $g_0$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b><math>0.6525 \pm 0.0009 \pm 0.0033</math></b>	60M	<sup>1</sup> BATLEY	09a NA48	$\pm$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.645 \pm 0.004 \pm 0.009$	23M	<sup>2</sup> BATLEY	06b NA48	$\pm$

<sup>1</sup> This fit is obtained with the CABIBBO 05 matrix element in the  $2\pi^0$  invariant mass squared range  $0.074094 < m_{2\pi^0}^2 < 0.104244 \text{ GeV}^2$ . Electromagnetic corrections and CHPT constraints for  $\pi\pi$  phase shifts ( $a_0$  and  $a_2$ ) have been used. Also measured ( $a_0 - a_2$ )  $m_{\pi^+} = 0.2646 \pm 0.0021 \pm 0.0023$ , where  $k_0$  was kept fixed in the fit at  $-0.0099$ .

<sup>2</sup> Superseded by BATLEY 09a. This fit is obtained with the CABIBBO 05 matrix element in the  $2\pi^0$  invariant mass squared range  $0.074 \text{ GeV}^2 < m_{2\pi^0}^2 < 0.097 \text{ GeV}^2$ , assuming  $k = 0$  (no term proportional to  $(s_2 - s_1)^2$ ) and excluding the kinematic region around the cusp ( $m_{2\pi^0}^2 = (2m_{\pi^+})^2 \pm 0.000525 \text{ GeV}^2$ ). Also  $\pi\pi$  phase shifts  $a_0$  and  $a_2$  are measured: ( $a_0 - a_2$ )  $m_{\pi^+} = 0.268 \pm 0.010 \pm 0.004 \pm 0.013$ (external) and  $a_2 m_{\pi^+} = -0.041 \pm 0.022 \pm 0.014$ .

**QUADRATIC COEFFICIENT  $h'$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b><math>-0.0433 \pm 0.0008 \pm 0.0026</math></b>	60M	<sup>1</sup> BATLEY	09a NA48	$\pm$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.047 \pm 0.012 \pm 0.011$	23M	<sup>2</sup> BATLEY	06b NA48	$\pm$

<sup>1</sup> This fit is obtained with the CABIBBO 05 matrix element in the  $2\pi^0$  invariant mass squared range  $0.074094 < m_{2\pi^0}^2 < 0.104244 \text{ GeV}^2$ . Electromagnetic corrections and CHPT constraints for  $\pi\pi$  phase shifts ( $a_0$  and  $a_2$ ) have been used. Also measured ( $a_0 - a_2$ )  $m_{\pi^+} = 0.2646 \pm 0.0021 \pm 0.0023$ , where  $k_0$  was kept fixed in the fit at  $-0.0099$ .

<sup>2</sup> Superseded by BATLEY 09a. This fit is obtained with the CABIBBO 05 matrix element in the  $2\pi^0$  invariant mass squared range  $0.074 \text{ GeV}^2 < m_{2\pi^0}^2 < 0.097 \text{ GeV}^2$ , assuming  $k = 0$  (no term proportional to  $(s_2 - s_1)^2$ ) and excluding the kinematic region around the cusp ( $m_{2\pi^0}^2 = (2m_{\pi^+})^2 \pm 0.000525 \text{ GeV}^2$ ). Also  $\pi\pi$  phase shifts  $a_0$  and  $a_2$  are measured: ( $a_0 - a_2$ )  $m_{\pi^+} = 0.268 \pm 0.010 \pm 0.004 \pm 0.013$ (external) and  $a_2 m_{\pi^+} = -0.041 \pm 0.022 \pm 0.014$ .

**QUADRATIC COEFFICIENT  $k_0$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b><math>0.0095 \pm 0.00017 \pm 0.00048</math></b>	60M	<sup>1</sup> BATLEY	09a NA48	$\pm$

<sup>1</sup> Assumed  $a_2 m_{\pi^+} = -0.0044$  in the fit.

**LINEAR COEFFICIENT  $g_{BB}$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b><math>0.6219 \pm 0.0009 \pm 0.0033</math></b>	60M	<sup>1</sup> BATLEY	09a NA48	$\pm$

<sup>1</sup> This fit is obtained using parametrizations of COLANGELO 06a and BISSEGGER 09 in the  $2\pi^0$  invariant mass squared range  $0.074094 < m_{2\pi^0}^2 < 0.104244 \text{ GeV}^2$ . Electromagnetic corrections and CHPT constraints for  $\pi\pi$  phase shifts ( $a_0$  and  $a_2$ ) have been used. Also measured ( $a_0 - a_2$ )  $m_{\pi^+} = 0.2633 \pm 0.0024 \pm 0.0024$ , where  $k_0$  was kept fixed in the fit at 0.0085.

**QUADRATIC COEFFICIENT  $h'_{BB}$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$** 

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b><math>-0.0520 \pm 0.0009 \pm 0.0026</math></b>	60M	<sup>1</sup> BATLEY	09a NA48	$\pm$

<sup>1</sup> This fit is obtained using parametrizations of COLANGELO 06a and BISSEGGER 09 in the  $2\pi^0$  invariant mass squared range  $0.074094 < m_{2\pi^0}^2 < 0.104244 \text{ GeV}^2$ . Electromagnetic corrections and CHPT constraints for  $\pi\pi$  phase shifts ( $a_0$  and  $a_2$ ) have been used. Also measured ( $a_0 - a_2$ )  $m_{\pi^+} = 0.2633 \pm 0.0024 \pm 0.0024$ , where  $k_0$  was kept fixed in the fit at 0.0085.

 **$K_{\ell 3}^\pm$  AND  $K_{\ell 3}^0$  FORM FACTORS**

Updated September 2013 by T.G. Trippe (LBNL) and C.-J. Lin (LBNL).

Assuming that only the vector current contributes to  $K \rightarrow \pi \ell \nu$  decays, we write the matrix element as

$$M \propto f_+(t) [(P_K + P_\pi)_\mu \bar{\ell} \gamma_\mu (1 + \gamma_5) \nu] + f_-(t) [m_\ell \bar{\ell} (1 + \gamma_5) \nu], \quad (1)$$

where  $P_K$  and  $P_\pi$  are the four-momenta of the  $K$  and  $\pi$  mesons,  $m_\ell$  is the lepton mass, and  $f_+$  and  $f_-$  are dimensionless form factors which can depend only on  $t = (P_K - P_\pi)^2$ , the square of the four-momentum transfer to the leptons. If time-reversal invariance holds,  $f_+$  and  $f_-$  are relatively real.  $K_{\mu 3}$  experiments, discussed immediately below, measure  $f_+$  and  $f_-$ , while  $K_{e 3}$  experiments, discussed further below, are sensitive only to  $f_+$  because the small electron mass makes the  $f_-$  term negligible.

**$K_{\mu 3}$  Experiments.** Analyses of  $K_{\mu 3}$  data frequently assume a linear dependence of  $f_+$  and  $f_-$  on  $t$ , i.e.,

$$f_\pm(t) = f_\pm(0) [1 + \lambda_\pm(t/m_{\pi^+}^2)]. \quad (2)$$

Most  $K_{\mu 3}$  data are adequately described by Eq. (2) for  $f_+$  and a constant  $f_-$  (i.e.,  $\lambda_- = 0$ ).

There are two equivalent parametrizations commonly used in these analyses:

**(1)  $\lambda_+, \xi(0)$  parametrization.** Older analyses of  $K_{\mu 3}$  data often introduce the ratio of the two form factors

$$\xi(t) = f_-(t)/f_+(t). \quad (3)$$

The  $K_{\mu 3}$  decay distribution is then described by the two parameters  $\lambda_+$  and  $\xi(0)$  (assuming time reversal invariance and  $\lambda_- = 0$ ).

**(2)  $\lambda_+, \lambda_0$  parametrization.** More recent  $K_{\mu 3}$  analyses have parametrized in terms of the form factors  $f_+$  and  $f_0$ , which are associated with vector and scalar exchange, respectively, to the lepton pair.  $f_0$  is related to  $f_+$  and  $f_-$  by

$$f_0(t) = f_+(t) + [t/(m_K^2 - m_\pi^2)] f_-(t). \quad (4)$$

# Meson Particle Listings

$K^\pm$

Here  $f_0(0)$  must equal  $f_+(0)$  unless  $f_-(t)$  diverges at  $t = 0$ . The earlier assumption that  $f_+$  is linear in  $t$  and  $f_-$  is constant leads to  $f_0$  linear in  $t$ :

$$f_0(t) = f_0(0) [1 + \lambda_0(t/m_{\pi^+}^2)] . \quad (5)$$

With the assumption that  $f_0(0) = f_+(0)$ , the two parametrizations,  $(\lambda_+, \xi(0))$  and  $(\lambda_+, \lambda_0)$  are equivalent as long as correlation information is retained.  $(\lambda_+, \lambda_0)$  correlations tend to be less strong than  $(\lambda_+, \xi(0))$  correlations.

Since the 2006 edition of the *Review* [4], we no longer quote results in the  $(\lambda_+, \xi(0))$  parametrization. We have removed many older low statistics results from the Listings. See the 2004 version of this note [5] for these older results, and the 1982 version [6] for additional discussion of the  $K_{\mu 3}^0$  parameters, correlations, and conversion between parametrizations.

**Quadratic Parametrization.** More recent high-statistics experiments have included a quadratic term in the expansion of  $f_+(t)$ ,

$$f_+(t) = f_+(0) \left[ 1 + \lambda'_+(t/m_{\pi^+}^2) + \frac{\lambda''_+}{2}(t/m_{\pi^+}^2)^2 \right] . \quad (6)$$

If there is a non-vanishing quadratic term, then  $\lambda_+$  of Eq. (2) represents the average slope, which is then different from  $\lambda'_+$ . Our convention is to include the factor  $\frac{1}{2}$  in the quadratic term, and to use  $m_{\pi^+}$  even for  $K_{e3}^+$  and  $K_{\mu 3}^+$  decays. We have converted other's parametrizations to match our conventions, as noted in the beginning of the “ $K_{e3}^\pm$  and  $K_{\mu 3}^0$  Form Factors” sections of the Listings.

**Pole Parametrization:** The pole model describes the  $t$ -dependence of  $f_+(t)$  and  $f_0(t)$  in terms of the exchange of the lightest vector and scalar  $K^*$  mesons with masses  $M_v$  and  $M_s$ , respectively:

$$f_+(t) = f_+(0) \left[ \frac{M_v^2}{M_v^2 - t} \right] , \quad f_0(t) = f_0(0) \left[ \frac{M_s^2}{M_s^2 - t} \right] . \quad (7)$$

**Dispersive Parametrization** [7,8]. This approach uses dispersive techniques and the known low-energy  $K$ - $\pi$  phases to parametrize the vector and scalar form factors:

$$f_+(t) = f_+(0) \exp \left[ \frac{t}{m_\pi^2} (\Lambda_+ + H(t)) \right] ; \quad (8)$$

$$f_0(t) = f_0(0) \exp \left[ \frac{t}{(m_K^2 - m_\pi^2)} (\ln[C] - G(t)) \right] , \quad (9)$$

where  $\Lambda_+$  is the slope of the vector form factor, and  $\ln[C] = \ln[f_0(m_K^2 - m_\pi^2)]$  is the logarithm of the scalar form factor at the Callan-Treiman point. The functions  $H(t)$  and  $G(t)$  are dispersive integrals.

**$K_{e3}$  Experiments:** Analysis of  $K_{e3}$  data is simpler than that of  $K_{\mu 3}$  because the second term of the matrix element assuming a pure vector current [Eq. (1) above] can be neglected. Here  $f_+$  can be assumed to be linear in  $t$ , in which case the linear coefficient  $\lambda_+$  of Eq. (2) is determined, or quadratic, in which case the linear coefficient  $\lambda'_+$  and quadratic coefficient  $\lambda''_+$  of Eq. (6) are determined.

If we remove the assumption of a pure vector current, then the matrix element for the decay, in addition to the terms in Eq. (1), would contain

$$+2m_K f_S \bar{\ell}(1 + \gamma_5)\nu \\ + (2f_T/m_K)(P_K)_\lambda (P_\pi)_\mu \bar{\ell} \sigma_{\lambda\mu} (1 + \gamma_5)\nu , \quad (10)$$

where  $f_S$  is the scalar form factor, and  $f_T$  is the tensor form factor. In the case of the  $K_{e3}$  decays where the  $f_-$  term can be neglected, experiments have yielded limits on  $|f_S/f_+|$  and  $|f_T/f_+|$ .

**Fits for  $K_{e3}$  Form Factors.** For  $K_{e3}$  data, we determine best values for the three parametrizations: linear  $(\lambda_+)$ , quadratic  $(\lambda'_+, \lambda''_+)$  and pole  $(M_v)$ . For  $K_{\mu 3}$  data, we determine best values for the three parametrizations: linear  $(\lambda_+, \lambda_0)$ , quadratic  $(\lambda'_+, \lambda''_+, \lambda_0)$  and pole  $(M_v, M_s)$ . We then assume  $\mu - e$  universality so that we can combine  $K_{e3}$  and  $K_{\mu 3}$  data, and again determine best values for the three parametrizations: linear  $(\lambda_+, \lambda_0)$ , quadratic  $(\lambda'_+, \lambda''_+, \lambda_0)$ , and pole  $(M_v, M_s)$ . When there is more than one parameter, fits are done including input correlations. Simple averages suffice in the two  $K_{e3}$  cases where there is only one parameter: linear  $(\lambda_+)$  and pole  $(M_v)$ .

Both KTeV and KLOE see an improvement in the quality of their fits relative to linear fits when a quadratic term is introduced, as well as when the pole parametrization is used. The quadratic parametrization has the disadvantage that the quadratic parameter  $\lambda''_+$  is highly correlated with the linear parameter  $\lambda'_+$ , in the neighborhood of 95%, and that neither parameter is very well determined. The pole fit has the same number of parameters as the linear fit, but yields slightly better fit probabilities, so that it would be advisable for all experiments to include the pole parametrization as one of their choices [9].

The “Kaon Particle Listings” show the results with and without assuming  $\mu - e$  universality. The “Meson Summary Tables” show all of the results assuming  $\mu - e$  universality, but most results not assuming  $\mu - e$  universality are given only in the Listings.

## References

1. L.M. Chounet, J.M. Gaillard, and M.K. Gaillard, Phys. Reports **4C**, 199 (1972).
2. H.W. Fearing, E. Fischbach, and J. Smith, Phys. Rev. **D2**, 542 (1970).
3. N. Cabibbo and A. Maksymowicz, Phys. Lett. **9**, 352 (1964).
4. W.-M. Yao *et al.*, Particle Data Group, J. Phys. **G33**, 1 (2006).
5. S. Eidelman *et al.*, Particle Data Group, Phys. Lett. **B592**, 1 (2004).
6. M. Roos *et al.*, Particle Data Group, Phys. Lett. **111B**, 73 (1982).
7. V. Bernard *et al.*, Phys. Lett. **B638**, 48 (2006).
8. A. Lai *et al.*, Phys. Lett. **B647**, 341 (2007), and references therein.
9. We thank P. Franzini (Rome U. and Frascati) for useful discussions on this point.

**$K_{e3}^\pm$  FORM FACTORS**

In the form factor comments, the following symbols are used.

$f_+$  and  $f_-$  are form factors for the vector matrix element.

$f_S$  and  $f_T$  refer to the scalar and tensor term.

$f_0 = f_+ + f_- t/(m_{K^+}^2 - m_{\pi^0}^2)$ .

$t$  = momentum transfer to the  $\pi$ .

$\lambda_+$  and  $\lambda_0$  are the linear expansion coefficients of  $f_+$  and  $f_0$ :

$f_+(t) = f_+(0) (1 + \lambda_+ t/m_{\pi^+}^2)$

For quadratic expansion

$f_+(t) = f_+(0) (1 + \lambda'_+ t/m_{\pi^+}^2 + \frac{\lambda''_+}{2} t^2/m_{\pi^+}^4)$

as used by KTeV. If there is a non-vanishing quadratic term, then  $\lambda_+$  represents an average slope, which is then different from  $\lambda'_+$ .

NA48 and ISTRA quadratic expansion coefficients are converted with  $\lambda'^+_{PDG} = \lambda_{NA48}$  and  $\lambda''_{PDG} = 2 \lambda'_{NA48}$

$\lambda'^+_{PDG} = (\frac{m_{\pi^+}}{m_{\pi^0}})^2 \lambda_+ ISTRA$  and

$\lambda''_{PDG} = 2 (\frac{m_{\pi^+}}{m_{\pi^0}})^4 \lambda'_+ ISTRA$

ISTRA linear expansion coefficients are converted with

$\lambda^+_{PDG} = (\frac{m_{\pi^+}}{m_{\pi^0}})^2 \lambda_+ ISTRA$  and  $\lambda_0_{PDG} = (\frac{m_{\pi^+}}{m_{\pi^0}})^2 \lambda_0 ISTRA$

The pole parametrization is

$f_+(t) = f_+(0) (\frac{M_V^2}{M_V^2 - t})$

$f_0(t) = f_0(0) (\frac{M_S^2}{M_S^2 - t})$

where  $M_V$  and  $M_S$  are the vector and scalar pole masses.

The following abbreviations are used:

DP = Dalitz plot analysis.

PI =  $\pi$  spectrum analysis.

MU =  $\mu$  spectrum analysis.

POL =  $\mu$  polarization analysis.

BR =  $K_{\mu 3}^\pm/K_{e 3}^\pm$  branching ratio analysis.

E = positron or electron spectrum analysis.

RC = radiative corrections.

 **$\lambda_+$  (LINEAR ENERGY DEPENDENCE OF  $f_+$  IN  $K_{e3}^\pm$  DECAY)**

These results are for a linear expansion only. See the next section for fits including a quadratic term. For radiative correction of the  $K_{e3}^\pm$  Dalitz plot, see GINSBERG 67, BECHERRAWY 70, CIRIGLIANO 02, CIRIGLIANO 04, and ANDRE 07. Results labeled OUR FIT are discussed in the review " $K_{e3}^\pm$  and  $K_{\mu 3}^0$  Form Factors" above. For earlier, lower statistics results, see the 2004 edition of this review, Physics Letters **B592** 1 (2004).

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.97 ± 0.05 OUR FIT</b>	Assuming $\mu$ -e universality				
<b>2.98 ± 0.05 OUR AVERAGE</b>					
3.044 ± 0.083 ± 0.074	1.1M	AKOPDZANOV 09	TNF	±	
2.966 ± 0.050 ± 0.034	919k	<sup>1</sup> YUSHCHENKO 04B	ISTR	—	DP
2.78 ± 0.26 ± 0.30	41k	SHIMIZU 00	SPEC	+	DP
2.84 ± 0.27 ± 0.20	32k	<sup>2</sup> AKIMENKO 91	SPEC	PI, no RC	
2.9 ± 0.4	62k	<sup>3</sup> BOLOTOV 88	SPEC	PI, no RC	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.06 ± 0.09 ± 0.06	550k	<sup>1,4</sup> AJINENKO 03c	ISTR	—	DP
2.93 ± 0.15 ± 0.2	130k	<sup>4</sup> AJINENKO 02	SPEC	DP	

<sup>1</sup> Rescaled to agree with our conventions as noted above.

<sup>2</sup> AKIMENKO 91 state that radiative corrections would raise  $\lambda_+$  by 0.0013.

<sup>3</sup> BOLOTOV 88 state radiative corrections of GINSBERG 67 would raise  $\lambda_+$  by 0.002.

<sup>4</sup> Superseded by YUSHCHENKO 04B.

 **$\lambda_+$  (LINEAR ENERGY DEPENDENCE OF  $f_+$  IN  $K_{\mu 3}^\pm$  DECAY)**

Results labeled OUR FIT are discussed in the review " $K_{e3}^\pm$  and  $K_{\mu 3}^0$  Form Factors" above. For earlier, lower statistics results, see the 2004 edition of this review, Physics Letters **B592** 1 (2004).

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.97 ± 0.05 OUR FIT</b>	Assuming $\mu$ -e universality				
<b>2.96 ± 0.17 OUR FIT</b>	Not assuming $\mu$ -e universality				
2.96 ± 0.14 ± 0.10	540k	<sup>1</sup> YUSHCHENKO04	ISTR	—	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.21 ± 0.45	112k	<sup>2</sup> AJINENKO 03	ISTR	—	DP

<sup>1</sup> Rescaled to agree with our conventions as noted above.

<sup>2</sup> Superseded by YUSHCHENKO 04.

 **$\lambda_0$  (LINEAR ENERGY DEPENDENCE OF  $f_0$  IN  $K_{e3}^\pm$  DECAY)**

Results labeled OUR FIT are discussed in the review " $K_{e3}^\pm$  and  $K_{\mu 3}^0$  Form Factors" above. For earlier, lower statistics results, see the 2004 edition of this review, Physics Letters **B592** 1 (2004).

VALUE (units $10^{-2}$ )	$d\lambda_0/d\lambda_+$	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.95 ± 0.12 OUR FIT</b>	Assuming $\mu$ -e universality					
<b>1.96 ± 0.13 OUR FIT</b>	Not assuming $\mu$ -e universality					
+1.96 ± 0.12 ± 0.06	−0.348	540k	<sup>1</sup> YUSHCHENKO04	ISTR	—	DP

• • • We do not use the following data for averages, fits, limits, etc. • • •

+2.09 ± 0.45 −0.46 112k <sup>2</sup>AJINENKO 03 ISTR — DP

+1.9 ± 0.64 24k <sup>3</sup>HORIE 01 SPEC + BR

+1.9 ± 1.0 +0.03 55k <sup>4</sup>HEINTZE 77 SPEC + BR

<sup>1</sup> Rescaled to agree with our conventions as noted above.

<sup>2</sup> Superseded by YUSHCHENKO 04.

<sup>3</sup> HORIE 01 assumes  $\mu$ -e universality in  $K_{e3}^\pm$  decay and uses SHIMIZU 00 value  $\lambda = 0.0278 \pm 0.0040$  from  $K_{e3}^\pm$  decay.

<sup>4</sup> HEINTZE 77 uses  $\lambda_+ = 0.029 \pm 0.003$ .  $d\lambda_0/d\lambda_+$  estimated by us.

 **$\lambda'_+$  (LINEAR  $K_{e3}^\pm$  FORM FACTOR FROM QUADRATIC FIT)**

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.485 ± 0.163 ± 0.034</b>	919k	<sup>1,2</sup> YUSHCHENKO04B	ISTR	—	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.07 ± 0.21	550k	<sup>1,3</sup> AJINENKO 03c	ISTR	—	DP

<sup>1</sup> Rescaled to agree with our conventions as noted above.

<sup>2</sup> YUSHCHENKO 04B  $\lambda'_+$  and  $\lambda''_+$  are strongly correlated with coefficient  $\rho(\lambda'_+, \lambda''_+) = -0.95$ .

<sup>3</sup> Superseded by YUSHCHENKO 04B.

 **$\lambda''_+$  (QUADRATIC  $K_{e3}^\pm$  FORM FACTOR)**

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.192 ± 0.062 ± 0.071</b>	919k	<sup>1,2</sup> YUSHCHENKO04B	ISTR	—	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •					
−0.5 ± 0.7 ± 1.5	550k	<sup>1,3</sup> AJINENKO 03c	ISTR	—	DP

<sup>1</sup> Rescaled to agree with our conventions as noted above.

<sup>2</sup> YUSHCHENKO 04B  $\lambda'_+$  and  $\lambda''_+$  are strongly correlated with coefficient  $\rho(\lambda'_+, \lambda''_+) = -0.95$ .

<sup>3</sup> Superseded by YUSHCHENKO 04B.

 **$|f_S/f_+|$  FOR  $K_{e3}^\pm$  DECAY**

Ratio of scalar to  $f_+$  couplings.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>−0.3 ± 0.8 −0.7 OUR AVERAGE</b>						
−0.37 ± 0.66 ± 0.41	919k	YUSHCHENKO04B	ISTR	—	$\lambda'_+, \lambda''_+, f_S$ fit	
0.2 ± 2.6 ± 1.4	41k	SHIMIZU 00	SPEC	+	$\lambda_+, f_S, f_T$ fit	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.2 ± 2.0 ± 0.3	550k	<sup>1</sup> AJINENKO 03c	ISTR	—	$\lambda_+, f_S, f_T$ fit	
−1.9 ± 2.5 ± 1.6	130k	<sup>1</sup> AJINENKO 02	SPEC	$\lambda_+, f_S$ fit		
7.0 ± 1.6 ± 1.6	32k	AKIMENKO 91	SPEC	$\lambda_+, f_S, f_T, \phi$ fit		
0 ± 10	2827	<sup>2</sup> BRAUN 75	HLBC	+		
< 13	90	CHIANG 72	OSPK	+		
14 ± 3 ± 4	2707	<sup>2</sup> STEINER 71	HLBC	+	$\lambda_+, f_S, f_T, \phi$ fit	
< 23	90	BOTTERILL 68c	ASPK			
< 18	90	BELLOTTI 67B	HLBC			
< 30	95	KALMUS 67	HLBC	+		

<sup>1</sup> Superseded by YUSHCHENKO 04B.

<sup>2</sup> Statistical errors only.

 **$|f_T/f_+|$  FOR  $K_{e3}^\pm$  DECAY**

Ratio of tensor to  $f_+$  couplings.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>−1.2 ± 2.3 OUR AVERAGE</b>						
−1.2 ± 2.1 ± 1.1	919k	YUSHCHENKO04B	ISTR	—	$\lambda'_+, \lambda''_+, f_T$ fit	
1 ± 1 ± 9	41k	SHIMIZU 00	SPEC	+	$\lambda_+, f_S, f_T$ fit	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
2.1 ± 6.4 ± 2.6	550k	<sup>1</sup> AJINENKO 03c	ISTR	—	$\lambda_+, f_S, f_T$ fit	
−4.5 ± 6.0 ± 5.7	130k	<sup>1</sup> AJINENKO 02	SPEC	$\lambda_+, f_T$ fit		
53 ± 9 ± 10	32k	AKIMENKO 91	SPEC	$\lambda_+, f_S, f_T, \phi$ fit		
7 ± 37	2827	<sup>2</sup> BRAUN 75	HLBC	+		
< 75	90	CHIANG 72	OSPK	+		
24 ± 16 ± 14	2707	<sup>2</sup> STEINER 71	HLBC	+	$\lambda_+, f_S, f_T, \phi$ fit	
< 58	90	BOTTERILL 68c	ASPK			
< 58	90	BELLOTTI 67B	HLBC			
< 110	95	KALMUS 67	HLBC	+		

<sup>1</sup> Superseded by YUSHCHENKO 04B.

<sup>2</sup> Statistical errors only.

 **$f_S/f_+$  FOR  $K_{\mu 3}^\pm$  DECAY**

Ratio of scalar to  $f_+$  couplings.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.17 ± 0.14 ± 0.54</b>	540k	<sup>1</sup> YUSHCHENKO04	ISTR	—	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.4 ± 0.5 ± 0.5	112k	<sup>2</sup> AJINENKO 03	ISTR	—	DP

<sup>1</sup> The second error is the theoretical error from the uncertainty in the chiral perturbation theory prediction for  $\lambda_0$ ,  $\pm 0.0053$ , combined in quadrature with the systematic error  $\pm 0.0009$ .

<sup>2</sup> The second error is the theoretical error from the uncertainty in the chiral perturbation theory prediction for  $\lambda_0$ . Superseded by YUSHCHENKO 04.

# Meson Particle Listings

## $K^\pm$

### $f_T/f_+$ FOR $K^\pm_{S3}$ DECAY

Ratio of tensor to  $f_+$  couplings.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$-0.07 \pm 0.71 \pm 0.20$	540k	YUSHCHENKO04	ISTR	—	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$-2.1 \pm 2.8 \pm 1.4$	112k	<sup>1</sup> AJINENKO	03	ISTR	— DP
$2 \pm 12$	1585	BRAUN	75	HLBC	

<sup>1</sup> The second error is the theoretical error from the uncertainty in the chiral perturbation theory prediction for  $\lambda_0$ . Superseded by YUSHCHENKO 04.

### $K_{\ell 4}^\pm$ FORM FACTORS

Based on the parametrizations of AMOROS 99, the  $K_{\ell 4}^\pm$  form factors can be expressed as

$$F_s = f_s + f'_s q^2 + f''_s q^4 + f'_e S_e / 4m_\pi^2$$
$$F_p = f_p$$
$$G_p = g_p + g'_p q^2$$
$$H_p = h_p$$

where  $q^2 = (S_\pi / 4m_\pi^2) - 1$ ,  $S_\pi$  is the invariant mass squared of the dipion, and  $S_e$  is the invariant mass squared of the dilepton.

### $f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>5.712±0.032 OUR AVERAGE</b>				
5.705±0.003±0.035	1.1M	<sup>1</sup> BATLEY	12	NA48 ±
5.75±0.02 ±0.08	400k	<sup>2</sup> PISLAK	03	B865 +

<sup>1</sup> BATLEY 12 uses data collected in 2003–2004. The result is obtained from a measurement of  $\Gamma(\pi^+ \pi^- e \nu)/\Gamma(\pi^+ \pi^- \pi^+)$  and assumed PDG 12 value of  $\Gamma(\pi^+ \pi^- \pi^+)/\Gamma = (5.59 \pm 0.04) \times 10^{-2}$ .

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, PISLAK 03 obtains the following  $\pi\pi$  scattering lengths  $a_0^0 = 0.228 \pm 0.012 \pm 0.004 \pm_{0.016}^{0.012}$  (theor.) and  $a_0^2 = -0.0365 \pm 0.0023 \pm 0.0008 \pm_{0.0026}^{0.0031}$  (theor.).

### $f'_s/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>15.2±0.7±0.5</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17.2±0.9±0.6	670k	<sup>2</sup> BATLEY	08A	NA48 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). The correlation with  $f'_s/f_s = -0.954$  and with  $f'_e/f_s = 0.080$ . Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

### $f''_s/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>-7.3±0.7±0.6</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-9.0±0.9±0.7	670k	<sup>2</sup> BATLEY	08A	NA48 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). The correlation with  $f'_s/f_s = -0.954$  and with  $f'_e/f_s = 0.019$ . Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

### $f'_e/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>6.8±0.6±0.7</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.1±0.8±0.9	670k	<sup>2</sup> BATLEY	08A	NA48 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). The correlation with  $f'_s/f_s = 0.080$  and with  $f''_s/f_s = 0.019$ . Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

### $f_p/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>-4.8±0.3±0.4</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-4.8±0.4±0.4	670k	<sup>2</sup> BATLEY	08A	NA48 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

### $g_p/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>86.8±1.0±1.0</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
87.3±1.3±1.2	670k	<sup>2</sup> BATLEY	08A	NA48 ±
80.9±0.9±1.2	400k	<sup>3</sup> PISLAK	03	B865 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). Supersedes BATLEY 08A. The correlation with  $g'_p/f_s = -0.914$ . Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

<sup>3</sup> Radiative corrections included. Using Roy equations PISLAK 03 obtains the following scattering lengths  $a_0^0 = 0.203 \pm 0.033 \pm 0.004$ ,  $a_0^2 = -0.055 \pm 0.023 \pm 0.003$ .

### $g'_p/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>8.9±1.7±1.3</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.1±2.2±1.5	670k	<sup>2</sup> BATLEY	08A	NA48 ±
12.0±1.9±0.7	400k	<sup>3</sup> PISLAK	03	B865 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). The correlation with  $g_p/f_s = -0.914$ . Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

<sup>3</sup> Radiative corrections included. Using Roy equations PISLAK 03 obtains the following scattering lengths  $a_0^0 = 0.203 \pm 0.033 \pm 0.004$ ,  $a_0^2 = -0.055 \pm 0.023 \pm 0.003$ .

### $h_p/f_s$ FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu$ DECAY

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>-39.8±1.5±0.8</b>	1.13M	<sup>1</sup> BATLEY	10c	NA48 ±
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-41.1±1.9±0.8	670k	<sup>2</sup> BATLEY	08A	NA48 ±
-51.3±3.3±3.5	400k	<sup>3</sup> PISLAK	03	B865 ±

<sup>1</sup> Radiative corrections included. Using Roy equations and including isospin breaking, BATLEY 10c obtains the following scattering lengths  $a_0^0 = 0.2220 \pm 0.0128 \pm 0.0050 \pm 0.0037$  (theor.),  $a_0^2 = -0.0432 \pm 0.0086 \pm 0.0034 \pm 0.0028$  (theor.). Supersedes BATLEY 08A.

<sup>2</sup> Radiative corrections included. Using Roy equations and not including isospin breaking, BATLEY 08A obtains the following  $\pi\pi$  scattering length  $a_0^0 = 0.233 \pm 0.016 \pm 0.007$   $a_0^2 = -0.0471 \pm 0.011 \pm 0.004$ .

<sup>3</sup> Radiative corrections included. Using Roy equations PISLAK 03 obtains the following scattering lengths  $a_0^0 = 0.203 \pm 0.033 \pm 0.004$ ,  $a_0^2 = -0.055 \pm 0.023 \pm 0.003$ .

### DECAY FORM FACTOR FOR $K^\pm \rightarrow \pi^0 \pi^0 e^\pm \nu$

Given in BOLOTOV 86b, BARMIN 88b, and SHIMIZU 04.

### $K^\pm \rightarrow \ell^\pm \nu \gamma$ FORM FACTORS

For definitions of the axial-vector  $F_A$  and vector  $F_V$  form factor, see the “Note on  $\pi^\pm \rightarrow \ell^\pm \nu \gamma$  and  $K^\pm \rightarrow \ell^\pm \nu \gamma$  Form Factors” in the  $\pi^\pm$  section. In the kaon literature, often different definitions  $a_K = F_A/m_K$  and  $v_K = F_V/m_K$  are used.

### $F_A + F_V$ , SUM OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR $K \rightarrow e \nu_e \gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.133±0.008 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
0.125±0.007±0.001	1.4K	<sup>1</sup> AMBROSINO	09E	KLOE $E_\gamma$ in 10–250 MeV, $p_e > 200$ MeV/c
0.147±0.011	51	<sup>2</sup> HEINTZE	79	SPEC
0.150 $\pm_{-0.023}^{+0.018}$	56	<sup>3</sup> HEARD	75	SPEC

See key on page 601

## Meson Particle Listings

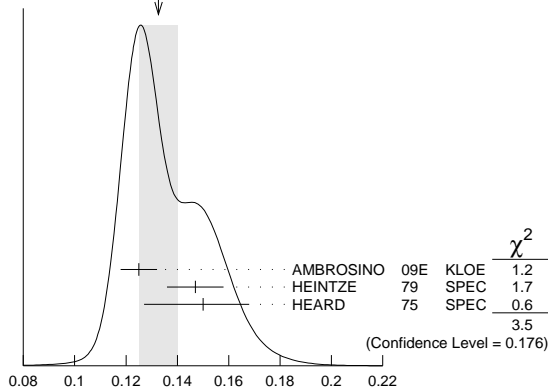
 $K^\pm$ 

<sup>1</sup> Vector form factor fitted with a linear function,  $V(x) = F_V (1 + \lambda(1-x))$ ,  $x = 2E_\gamma/m_K$ . The fitted value of  $\lambda = 0.38 \pm 0.20 \pm 0.02$  with a correlation of  $-0.93$  between  $(F_V + F_A)$  and  $\lambda$ .

<sup>2</sup> HEINTZE 79 quotes absolute value of  $|F_A + F_V| \sin\theta_C$ . We use  $\sin\theta_C = V_{us} = 0.2205$ .

<sup>3</sup> HEARD 75 quotes absolute value of  $|F_A + F_V| \sin\theta_C$ . We use  $\sin\theta_C = V_{us} = 0.2205$ .

WEIGHTED AVERAGE  
0.133±0.008 (Error scaled by 1.3)



$F_A + F_V$ , SUM OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR  $K \rightarrow e \nu \gamma$

$F_A + F_V$ , SUM OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR  $K \rightarrow \mu \nu \mu \gamma$

VALUE	CL%	EVTs	DOCUMENT ID	TECN	CHG
$0.165 \pm 0.007 \pm 0.011$		2588	<sup>1</sup> ADLER 00B	B787	+
• • • We do not use the following data for averages, fits, limits, etc. • • •					
-1.2 to 1.1	90		DEMIDOV 90	XEBC	
< 0.23	90		<sup>1</sup> AKIBA 85	SPEC	

<sup>1</sup> Quotes absolute value. Sign not determined.

$F_A - F_V$ , DIFFERENCE OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR  $K \rightarrow e \nu \gamma$

VALUE	CL%	DOCUMENT ID	TECN
$< 0.49$	90	<sup>1</sup> HEINTZE 79	SPEC

<sup>1</sup> HEINTZE 79 quotes  $|F_A - F_V| < \sqrt{11} |F_A + F_V|$ .

$F_A - F_V$ , DIFFERENCE OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR  $K \rightarrow \mu \nu \mu \gamma$

VALUE	CL%	EVTs	DOCUMENT ID	TECN	CHG
$-0.24$ to $0.04$	90	2588	ADLER 00B	B787	+
• • • We do not use the following data for averages, fits, limits, etc. • • •					
-2.2 to 0.6	90		DEMIDOV 90	XEBC	
-2.5 to 0.3	90		AKIBA 85	SPEC	

 $K^\pm$  CHARGE RADIUS

VALUE (fm)	DOCUMENT ID	COMMENT
$0.560 \pm 0.031$ OUR AVERAGE		
$0.580 \pm 0.040$	AMENDOLIA 86B	$K_e \rightarrow K_e$
$0.530 \pm 0.050$	DALLY 80	$K_e \rightarrow K_e$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$0.620 \pm 0.037$	BLATNIK 79	VMD + dispersion relations

 $K^+$  LONGITUDINAL POLARIZATION OF EMITTED  $\mu^+$ 

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
$< -0.990$	90	<sup>1</sup> AOKI 94	SPEC	+	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< -0.990$	90	IMAZATO 92	SPEC	+	Repl. by AOKI 94
$-0.970 \pm 0.047$		2 YAMANAKA 86	SPEC	+	
$-1.0 \pm 0.1$		2 CUTTS 69	SPRK	+	
$-0.96 \pm 0.12$		2 COOMBES 57	CNTR	+	

<sup>1</sup> AOKI 94 measures  $\xi P_\mu = -0.996 \pm 0.0030 \pm 0.0048$ . The above limit is obtained by summing the statistical and systematic errors in quadrature, normalizing to the physically significant region ( $|\xi P_\mu| < 1$ ) and assuming that  $\xi=1$ , its maximum value.

<sup>2</sup> Assumes  $\xi=1$ .

FORWARD-BACKWARD ASYMMETRY IN  $K^\pm$  DECAYS

$$A_{FB}(K^\pm_{\pi\mu\mu}) = \frac{\Gamma(\cos\theta_{K\mu} > 0) - \Gamma(\cos\theta_{K\mu} < 0)}{\Gamma(\cos\theta_{K\mu} > 0) + \Gamma(\cos\theta_{K\mu} < 0)}$$

VALUE	CL%	DOCUMENT ID	TECN
$< 2.3 \times 10^{-2}$	90	<sup>1</sup> BATLEY 11A	NA48

<sup>1</sup> BATLEY 11A gives a corresponding value of the asymmetry  $A_{FB} = (-2.4 \pm 1.8) \times 10^{-2}$ .

 $K^\pm$  REFERENCES

BABUSCI 14B	PL B738 128	D. Babusci et al.	(KLOE, KLOE-2 Collab.)
BATLEY 14	PL B730 141	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 14A	JHEP 1408 159	J.R. Batley et al.	(CERN NA48/2 Collab.)
LAZZERONI 14	PL B732 65	C. Lazzeroni et al.	(CERN NA62 Collab.)
UVAROV 14	PAN 77 725	V. A. Uvarov et al.	(ISTRA+ Collab.)
Translated from YAF 77 765.			
LAZZERONI 13	PL B719 326	C. Lazzeroni et al.	(CERN NA62 Collab.)
BATLEY 12	PL B715 105	J.R. Batley et al.	(CERN NA48/2 Collab.)
PDG 12	PR D86 010001	J. Beringer et al.	(PDG Collab.)
BATLEY 11A	PL B697 107	J.R. Batley et al.	(CERN NA48/2 Collab.)
LAZZERONI 11	PL B698 105	C. Lazzeroni et al.	(CERN NA62 Collab.)
ADLER 10	PR D81 092001	S. Adler et al.	(BNL E787 Collab.)
BATLEY 10A	EPJ C68 75	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 10C	EPJ C70 635	J.R. Batley et al.	(CERN NA48/2 Collab.)
PDG 10	JP G37 075021	K. Nakamura et al.	(PDG Collab.)
PISLAK 10A	PRL 105 019901E	S. Pislak et al.	(BNL E865 Collab.)
AKOPDZANOV 09	PAN 71 2074	G.A. Akopdzhanyan et al.	(IHEP)
Translated from YAF 71 2108.			
AMBROSINO 09E	EPJ C64 627	F. Ambrosino et al.	(KLOE Collab.)
Also	EPJ C65 703 (erratum)	F. Ambrosino et al.	(KLOE Collab.)
BATLEY 09	PL B677 246	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 09A	EPJ C64 589	J.R. Batley et al.	(CERN NA48/2 Collab.)
BISSEGGGER 09	NP B806 178	M. Bissegger et al.	
AMBROSINO 08A	JHEP 0801 073	F. Ambrosino et al.	(KLOE Collab.)
AMBROSINO 08B	JHEP 0802 098	F. Ambrosino et al.	(KLOE Collab.)
AMBROSINO 08E	PL B666 305	F. Ambrosino et al.	(KLOE Collab.)
ARTAMONOV 08	PRL 101 191802	A.V. Artamonov et al.	(BNL E949 Collab.)
Also	PR D79 092004	A.V. Artamonov et al.	(BNL E949 Collab.)
BATLEY 08	PL B659 493	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 08A	EPJ C54 411	J.R. Batley et al.	(CERN NA48/2 Collab.)
AKIMENKO 07	PAN 70 702	S.A. Akimenko et al.	(ISTRA+ Collab.)
Translated from YAF 70 734.			
ANDRE 07	ANP 322 2518	T. Andre	(EFI)
BATLEY 07A	EPJ C50 329	J.R. Batley et al.	(CERN NA48/2 Collab.)
Also	EPJ C52 1021 (erratum)	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 07B	PL B649 349	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 07E	EPJ C52 875	J.R. Batley et al.	(CERN NA48/2 Collab.)
TCHIKILEV 07	PAN 70 29	O.G. Tchikilev et al.	(ISTRA+ Collab.)
ALIEV 06	EPJ C46 61	M.A. Aliev et al.	(KEK E470 Collab.)
AMBROSINO 06A	PL B632 76	F. Ambrosino et al.	(KLOE Collab.)
BATLEY 06	PL B634 474	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 06A	PL B638 22	J.R. Batley et al.	(CERN NA48/2 Collab.)
Also	PL B640 297 (erratum)	J.R. Batley et al.	(CERN NA48/2 Collab.)
BATLEY 06B	PL B633 173	J.R. Batley et al.	(CERN NA48/2 Collab.)
COLANGELO 06A	PL B638 187	G. Colangelo et al.	
MA 06	PL D73 037101	H. Ma et al.	(BNL E865 Collab.)
PDG 06	PL G33 1	W.-M. Yao et al.	(PDG Collab.)
SHIMIZU 06	PL B633 190	S. Shimizu et al.	(KEK E470 Collab.)
UVAROV 06	PAN 69 26	V.A. Uvarov et al.	(ISTRA+ Collab.)
AKOPDZHAN...05	EPJ C40 343	G.A. Akopdzhanyan et al.	(IHEP)
Also	PAN 68 948	G.A. Akopdzhanyan et al.	(IHEP)
Translated from YAF 68 986.			
AKOPDZHAN...05B	JETPL 82 675	G.A. Akopdzhanyan et al.	(IHEP)
Translated from ZETFP 82 771.			
ARTAMONOV 05	PL B623 192	A.V. Artamonov et al.	(BNL E949 Collab.)
CABIBBO 05	JHEP 0503 021	N. Cabibbo, G. Isidori	(CERN, ROMA1, FRAS)
SHER 05	PR D72 012005	A. Sher et al.	(BNL E865 Collab.)
ABE 04F	PRL 93 131601	M. Abe et al.	(KEK E246 Collab.)
Also	PR D73 072005	M. Abe et al.	(KEK E246 Collab.)
ADLER 04A	PR D70 037102	S. Adler et al.	(BNL E787 Collab.)
ALOISIO 04	PL B597 139	A. Aloisio et al.	(KLOE Collab.)
ANISIMOVSK...04	PRL 93 031801	V.V. Anisimovskiy et al.	(BNL E949 Collab.)
Also	PR D77 052003	S. Adler et al.	(BNL E949 Collab.)
CABIBBO 04A	PRL 93 121801	N. Cabibbo	(CERN, ROMA1)
CIRIGLIANO 04	PL C53 53	V. Cirigliano, H. Neufeld, H. Pichl	(CIT, VALLÉE)
PDG 04	PL B592 1	S. Edelmann et al.	(PDG Collab.)
SHIMIZU 04	PR D70 037101	S. Shimizu et al.	(KEK E470 Collab.)
YUSHCHENKO 04	PL B581 31	O.P. Yushchenko et al.	(INRM, INRM)
YUSHCHENKO 04B	PL B589 111	O.P. Yushchenko et al.	(INRM)
AJINENKO 03	PAN 66 105	I.V. Ajinenko et al.	(IHEP, INRM)
Translated from YAF 66 107.			
AJINENKO 03B	PL B567 159	I.V. Ajinenko et al.	(IHEP, INRM)
AJINENKO 03C	PL B574 14	I.V. Ajinenko et al.	(IHEP, INRM)
ALIEV 03	PL B554 7	M.A. Aliev et al.	(KEK E470 Collab.)
ANISIMOVSK...03	PL B562 166	V.V. Anisimovskiy et al.	
PISLAK 03	PR D67 072004	S. Pislak et al.	(BNL E865 Collab.)
Also	PR D81 119903E	S. Pislak et al.	(BNL E865 Collab.)
SHER 03	PRL 91 261802	A. Sher et al.	(BNL E865 Collab.)
ADLER 02	PRL 88 041803	S. Adler et al.	(BNL E787 Collab.)
ADLER 02B	PR D65 052009	S. Adler et al.	(BNL E787 Collab.)
ADLER 02C	PL B537 211	S. Adler et al.	(BNL E787 Collab.)
AJINENKO 02	PAN 65 2064	I.V. Ajinenko et al.	(IHEP, INRM)
Translated from YAF 65 2125.			
CIRIGLIANO 02	EPJ C23 121	V. Cirigliano et al.	(VIEN, VALE, MARS)
PARK 02	PRL 88 111801	H.K. Park et al.	(FNAL HyperC Collab.)
PDG 02	PR D66 010001	K. Hagiwara et al.	(PDG Collab.)
POBLAGUEV 02	PRL 89 061803	A.A. Poblaguev et al.	(BNL 865 Collab.)
ADLER 01	PR D63 032004	S. Adler et al.	(BNL E787 Collab.)
HORIE 01	PL B513 311	K. Horie et al.	(KEK E426 Collab.)
PISLAK 01	PRL 87 221801	S. Pislak et al.	(BNL E865 Collab.)
Also	PR D67 072004	S. Pislak et al.	(BNL E865 Collab.)
Also	PRL 105 019901E	S. Pislak et al.	(BNL E865 Collab.)
ADLER 00	PRL 84 3768	S. Adler et al.	(BNL E787 Collab.)
ADLER 00B	PRL 85 2256	S. Adler et al.	(BNL E787 Collab.)
ADLER 00C	PRL 85 4056	S. Adler et al.	(BNL E787 Collab.)
APPEL 00	PRL 85 2450	R. Appel et al.	(BNL E787 Collab.)
Also	Thesis, Yale Univ.	D.R. Bergman	
Also	Thesis, Univ. Zurich	S. Pislak	
APPEL 00B	PRL 85 2877	R. Appel et al.	(BNL 865 Collab.)
MA 00	PRL 84 2580	H. Ma et al.	(BNL 865 Collab.)
PDG 00	EPJ C15 1	D.E. Groom et al.	(PDG Collab.)
SHIMIZU 00	PL B495 33	S. Shimizu et al.	(KEK E246 Collab.)
ABE 99S	PRL 83 4253	M. Abe et al.	(KEK E246 Collab.)
AMOROS 99	JP G25 1607	G. Amoros, J. Bijnens	(LUND, HELS)
APPEL 99	PRL 83 4482	R. Appel et al.	(BNL 865 Collab.)
ADLER 98	PR D58 012003	S. Adler et al.	(BNL E787 Collab.)
BATUSOV 98	NP B516 3	V.Y. Batusov et al.	
DAMBROSIO 98A	JHEP 9808 004	G. D'Ambrosio et al.	
ADLER 97	PRL 79 2204	S. Adler et al.	(BNL E787 Collab.)
ADLER 97C	PRL 79 4756	S. Adler et al.	(BNL E787 Collab.)
BERGMAN 97	Thesis, Yale Univ.	D.R. Bergman	
KITCHING 97	PRL 79 4079	P. Kitching et al.	(BNL E787 Collab.)
PISLAK 97	Thesis, Univ. Zurich	S. Pislak	
ADLER 96	PRL 76 1421	S. Adler et al.	(BNL E787 Collab.)
KOPTYEV 95	JETPL 61 877	V.P. Koptev et al.	(PNPI)
Translated from ZETFP 61 865.			
AOKI 94	PR D50 69	M. Aoki et al.	(INUS, KEK, TOKMS)
ATIYA 93	PRL 70 2521	M.S. Atiya et al.	(BNL E787 Collab.)
Also	PRL 71 305 (erratum)	M.S. Atiya et al.	(BNL E787 Collab.)
ATIYA 93B	PR D48 R1	M.S. Atiya et al.	(BNL E787 Collab.)
ALLIEGRO 92	PRL 68 278	C. Alliegro et al.	(BNL, FNAL, PSI+)



$-0.087 \pm 0.046$	BLATNIK	79	VMD + dispersion relations
$-0.050 \pm 0.130$	FOETH	698	$K_S$ regen. by electrons

**T-VIOLATION PARAMETER IN  $K^0\text{-}\bar{K}^0$  MIXING**

The asymmetry  $A_T = \frac{\Gamma(\bar{K}^0 \rightarrow K^0) - \Gamma(K^0 \rightarrow \bar{K}^0)}{\Gamma(\bar{K}^0 \rightarrow K^0) + \Gamma(K^0 \rightarrow \bar{K}^0)}$  must vanish if  $T$ -invariance holds.

**ASYMMETRY  $A_T$  IN  $K^0\text{-}\bar{K}^0$  MIXING**

VALUE (units $10^{-3}$ )	EVTs	DOCUMENT ID	TECN.
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**$6.6 \pm 1.3 \pm 1.0$**  640k 1 ANGELOPO... 98E CPLR

<sup>1</sup> ANGELOPOULOS 98E measures the asymmetry  $A_T = [\Gamma(\bar{K}^0_{t=0} \rightarrow e^+ \pi^- \nu_{t=\tau}) - \Gamma(K^0_{t=0} \rightarrow e^- \pi^+ \bar{\nu}_{t=\tau})] / [\Gamma(\bar{K}^0_{t=0} \rightarrow e^+ \pi^- \nu_{t=\tau}) + \Gamma(K^0_{t=0} \rightarrow e^- \pi^+ \bar{\nu}_{t=\tau})]$  as a function of the neutral-kaon eigentime  $\tau$ . The initial strangeness of the neutral kaon is tagged by the charge of the accompanying charged kaon in the reactions  $p\bar{p} \rightarrow K^- \pi^+ K^0$  and  $p\bar{p} \rightarrow K^+ \pi^- \bar{K}^0$ . The strangeness at the time of the decay is tagged by the lepton charge. The reported result is the average value of  $A_T$  over the interval  $1\tau_S < \tau < 20\tau_S$ . From this value of  $A_T$  ANGELOPOULOS 01B, assuming  $CPT$  invariance in the  $e\pi\nu$  decay amplitude, determine the  $T$ -violating as  $\Delta S = \Delta S$  conserving parameter (for its definition, see Review below)  $4\text{Re}(\epsilon) = (6.2 \pm 1.4 \pm 1.0) \times 10^{-3}$ .

 **$CPT$  INVARIANCE TESTS IN NEUTRAL KAON DECAY**

Updated October 2013 by M. Antonelli (LNF-INFN, Frascati) and G. D'Ambrosio (INFN Sezione di Napoli).

$CPT$  theorem is based on three assumptions: quantum field theory, locality, and Lorentz invariance, and thus it is a fundamental probe of our basic understanding of particle physics. Strangeness oscillation in  $K^0 - \bar{K}^0$  system, described by the equation

$$i \frac{d}{dt} \begin{bmatrix} K^0 \\ \bar{K}^0 \end{bmatrix} = [M - i\Gamma/2] \begin{bmatrix} K^0 \\ \bar{K}^0 \end{bmatrix},$$

where  $M$  and  $\Gamma$  are hermitian matrices (see PDG review [1], references [2,3], and KLOE paper [5] for notations and previous literature), allows a very accurate test of  $CPT$  symmetry; indeed since  $CPT$  requires  $M_{11} = M_{22}$  and  $\Gamma_{11} = \Gamma_{22}$ , the mass and width eigenstates,  $K_{S,L}$ , have a  $CPT$ -violating piece,  $\delta$ , in addition to the usual  $CPT$ -conserving parameter  $\epsilon$ :

$$K_{S,L} = \frac{1}{\sqrt{2(1 + |\epsilon_{S,L}|^2)}} \left[ (1 + \epsilon_{S,L}) K^0 \pm (1 - \epsilon_{S,L}) \bar{K}^0 \right]$$

$$\epsilon_{S,L} = \frac{-i\Im(M_{12}) - \frac{1}{2}\Im(\Gamma_{12}) \mp \frac{1}{2} \left[ M_{11} - M_{22} - \frac{i}{2}(\Gamma_{11} - \Gamma_{22}) \right]}{m_L - m_S + i(\Gamma_S - \Gamma_L)/2}$$

$$\equiv \epsilon \pm \delta. \quad (1)$$

Using the phase convention  $\Im(\Gamma_{12}) = 0$ , we determine the phase of  $\epsilon$  to be  $\varphi_{SW} \equiv \arctan \frac{2(m_L - m_S)}{\Gamma_S - \Gamma_L}$ . Imposing unitarity to an arbitrary combination of  $K^0$  and  $\bar{K}^0$  wave functions, we obtain the Bell-Steinberger relation [4] connecting  $CP$  and  $CPT$  violation in the mass matrix to  $CP$  and  $CPT$  violation in the decay; in fact, neglecting  $\mathcal{O}(\epsilon)$  corrections to the coefficient of the  $CPT$ -violating parameter,  $\delta$ , we can write [5]

$$\left[ \frac{\Gamma_S + \Gamma_L}{\Gamma_S - \Gamma_L} + i \tan \phi_{SW} \right] \left[ \frac{\Re(\epsilon)}{1 + |\epsilon|^2} - i\Im(\delta) \right] =$$

$$\frac{1}{\Gamma_S - \Gamma_L} \sum_f A_L(f) A_S^*(f), \quad (2)$$

**Table 1:** Values, errors, and correlation coefficients for  $\Re(\delta)$ ,  $\Im(\delta)$ ,  $\Re(x_-)$ ,  $\Im(x_+)$ , and  $A_S + A_L$  obtained from a combined fit, including KLOE [5] and CPLEAR [14].

	value	Correlations coefficients			
$\Re(\delta)$	$(3.0 \pm 2.3) \times 10^{-4}$	1			
$\Im(\delta)$	$(-0.66 \pm 0.65) \times 10^{-2}$	-0.21	1		
$\Re(x_-)$	$(-0.30 \pm 0.21) \times 10^{-2}$	-0.21	-0.60	1	
$\Im(x_+)$	$(0.02 \pm 0.22) \times 10^{-2}$	-0.38	-0.14	0.47	1
$A_S + A_L$	$(-0.40 \pm 0.83) \times 10^{-2}$	-0.10	-0.63	0.99	0.43 1

where  $A_{L,S}(f) \equiv A(K_{L,S} \rightarrow f)$ . We stress that this relation is phase-convention-independent. The advantage of the neutral kaon system is that only a few decay modes give significant contributions to the r.h.s. in Eq. (2); in fact, defining for the hadronic modes

$$\alpha_i \equiv \frac{1}{\Gamma_S} \langle \mathcal{A}_L(i) \mathcal{A}_S^*(i) \rangle = \eta_i \mathcal{B}(K_S \rightarrow i),$$

$$i = \pi^0 \pi^0, \pi^+ \pi^- (\gamma), 3\pi^0, \pi^0 \pi^+ \pi^- (\gamma), \quad (3)$$

the recent data from CPLEAR, KLOE, KTeV, and NA48 have led to the following determinations (the analysis described in Ref. 5 has been updated by using the recent measurements of  $K_L$  branching ratios from KTeV [6,7], NA48 [8,9], and the results described in the  $CP$  violation in  $K_L$  decays minireview, and the recent KLOE result [10])

$$\alpha_{\pi^+ \pi^-} = ((1.112 \pm 0.010) + i(1.061 \pm 0.010)) \times 10^{-3},$$

$$\alpha_{\pi^0 \pi^0} = ((0.493 \pm 0.005) + i(0.471 \pm 0.005)) \times 10^{-3},$$

$$\alpha_{\pi^+ \pi^- \pi^0} = ((0 \pm 2) + i(0 \pm 2)) \times 10^{-6},$$

$$|\alpha_{\pi^0 \pi^0 \pi^0}| < 1.5 \times 10^{-6} \quad \text{at } 95\% \text{ CL}. \quad (4)$$

The semileptonic contribution to the right-handed side of Eq. (2) requires the determination of several observables: we define [2,3]

$$\mathcal{A}(K^0 \rightarrow \pi^- l^+ \nu) = \mathcal{A}_0(1 - y),$$

$$\mathcal{A}(K^0 \rightarrow \pi^+ l^- \nu) = \mathcal{A}_0^*(1 + y^*)(x_+ - x_-)^*,$$

$$\mathcal{A}(\bar{K}^0 \rightarrow \pi^+ l^- \nu) = \mathcal{A}_0^*(1 + y^*),$$

$$\mathcal{A}(\bar{K}^0 \rightarrow \pi^- l^+ \nu) = \mathcal{A}_0(1 - y)(x_+ + x_-), \quad (5)$$

where  $x_+$  ( $x_-$ ) describes the violation of the  $\Delta S = \Delta Q$  rule in  $CPT$ -conserving (violating) decay amplitudes, and  $y$  parametrizes  $CPT$  violation for  $\Delta S = \Delta Q$  transitions. Taking advantage of their tagged  $K^0(\bar{K}^0)$  beams, CPLEAR has measured  $\Im(x_+)$ ,  $\Re(x_-)$ ,  $\Im(\delta)$ , and  $\Re(\delta)$  [11]. These determinations have been improved in Ref. 5 by including the information  $A_S - A_L = 4[\Re(\delta) + \Re(x_-)]$ , where  $A_{L,S}$  are the  $K_L$  and  $K_S$  semileptonic charge asymmetries, respectively, from the PDG [12] and KLOE [13]. Here we are also including the  $T$ -violating asymmetry measurement from CPLEAR [14].

# Meson Particle Listings

$K^0$

The value  $A_S + A_L$  in Table 1 can be directly included in the semileptonic contributions to the Bell Steinberger relations in Eq. (2)

$$\begin{aligned} \sum_{\pi\ell\nu} \langle \mathcal{A}_L(\pi\ell\nu) \mathcal{A}_S^*(\pi\ell\nu) \rangle \\ = 2\Gamma(K_L \rightarrow \pi\ell\nu) (\Re(\epsilon) - \Re(y) - i(\Im(x_+) + \Im(\delta))) \\ = 2\Gamma(K_L \rightarrow \pi\ell\nu) ((A_S + A_L)/4 - i(\Im(x_+) + \Im(\delta))) . \end{aligned} \quad (6)$$

Defining

$$\alpha_{\pi\ell\nu} \equiv \frac{1}{\Gamma_S} \sum_{\pi\ell\nu} \langle \mathcal{A}_L(\pi\ell\nu) \mathcal{A}_S^*(\pi\ell\nu) \rangle + 2i \frac{\tau_{KS}}{\tau_{KL}} \mathcal{B}(K_L \rightarrow \pi\ell\nu) \Im(\delta) , \quad (7)$$

we find:

$$\alpha_{\pi\ell\nu} = ((-0.2 \pm 0.5) + i(0.1 \pm 0.5)) \times 10^{-5} .$$

Inserting the values of the  $\alpha$  parameters into Eq. (2), we find

$$\begin{aligned} \Re(\epsilon) &= (161.1 \pm 0.5) \times 10^{-5} , \\ \Im(\delta) &= (-0.7 \pm 1.4) \times 10^{-5} . \end{aligned} \quad (8)$$

The complete information on Eq. (8) is given in Table 2.

**Table 2:** Summary of results: values, errors, and correlation coefficients for  $\Re(\epsilon)$ ,  $\Im(\delta)$ ,  $\Re(\delta)$ , and  $\Re(x_-)$ .

	value	Correlations coefficients			
$\Re(\epsilon)$	$(161.1 \pm 0.5) \times 10^{-5}$	+1			
$\Im(\delta)$	$(-0.7 \pm 1.4) \times 10^{-5}$	+0.09	1		
$\Re(\delta)$	$(2.4 \pm 2.3) \times 10^{-4}$	+0.08	-0.12	1	
$\Re(x_-)$	$(-4.1 \pm 1.7) \times 10^{-3}$	+0.14	0.22	-0.43	1

Now the agreement with  $CPT$  conservation,  $\Im(\delta) = \Re(\delta) = \Re(x_-) = 0$ , is at 18% C.L.

The allowed region in the  $\Re(\epsilon) - \Im(\delta)$  plane at 68% CL and 95% C.L. is shown in the top panel of Fig. 1.

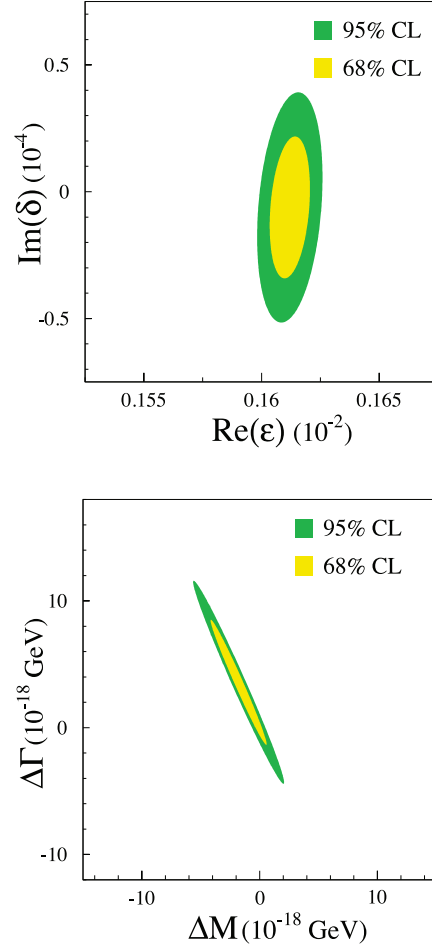
The process giving the largest contribution to the size of the allowed region is  $K_L \rightarrow \pi^+\pi^-$ , through the uncertainty on  $\phi_{+-}$ .

The limits on  $\Im(\delta)$  and  $\Re(\delta)$  can be used to constrain the  $K^0 - \bar{K}^0$  mass and width difference

$$\delta = \frac{i(m_{K^0} - m_{\bar{K}^0}) + \frac{1}{2}(\Gamma_{K^0} - \Gamma_{\bar{K}^0})}{\Gamma_S - \Gamma_L} \cos \phi_{SW} e^{i\phi_{SW}} [1 + \mathcal{O}(\epsilon)] .$$

The allowed region in the  $\Delta M = (m_{K^0} - m_{\bar{K}^0})$ ,  $\Delta\Gamma = (\Gamma_{K^0} - \Gamma_{\bar{K}^0})$  plane is shown in the bottom panel of Fig. 1. As a result, we improve on the previous limits (see for instance, P. Bloch in Ref. 12) and in the limit  $\Gamma_{K^0} - \Gamma_{\bar{K}^0} = 0$  we obtain

$$\begin{aligned} -4.0 \times 10^{-19} \text{ GeV} &< m_{K^0} - m_{\bar{K}^0} \\ &< 4.0 \times 10^{-19} \text{ GeV} \quad \text{at } 95 \% \text{ C.L.} \end{aligned}$$



**Figure 1:** Top: allowed region at 68% and 95% C.L. in the  $\Re(\epsilon)$ ,  $\Im(\delta)$  plane. Bottom: allowed region at 68% and 95% C.L. in the  $\Delta M$ ,  $\Delta\Gamma$  plane.

## References

1. See the “ $CP$  Violation in Meson Decays,” in this *Review*.
2. L. Maiani, “ $CP$  And  $CPT$  Violation in Neutral Kaon Decays,” L. Maiani, G. Pancheri, and N. Paver, *The Second Daphne Physics Handbook*, Vol. 1, 2.
3. G. D’Ambrosio, G. Isidori, and A. Pugliese, “ $CP$  and  $CPT$  measurements at DAΦNE,” L. Maiani, G. Pancheri, and N. Paver, *The Second Daphne Physics Handbook*, Vol. 1, 2.
4. J. S. Bell and J. Steinberger, In Wolfenstein, L. (ed.): *CP violation*, 42-57. (In *Oxford International Symposium Conference on Elementary Particles*, September 1965, 195-208, 221-222). (See Book Index).
5. F. Ambrosino *et al.*, [KLOE Collab.], JHEP **0612**, 011 (2006) [arXiv:hep-ex/0610034].
6. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. **D70**, 092006 (1998).
7. E. Abouzaid *et al.* [KTeV Collab.], Phys. Rev. **D83**, 092001 (2011).
8. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B645**, 26 (2007); A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B602**, 41 (2004).



9. We thank G. Isidori and M. Palutan for their contribution to the original analysis [5] performed with KLOE data.
10. D. Babusci *et al.*, [KLOE Collab.], Phys. Lett. **B723**, 54 (2013).
11. A. Angelopoulos *et al.*, [CPLEAR Collab.], Phys. Lett. **B444**, 52 (1998).
12. W. M. Yao *et al.*, [Particle Data Group], J. Phys. **G33**, 1 (2006).
13. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B636**, 173 (2006) [[arXiv:hep-ex/0601026](#)].
14. P. Bloch, M. Fidecaro, private communication of the data in a finer binning format; A. Angelopoulos *et al.*, [CPLEAR Collab.], Phys. Lett. **B444**, 43 (1998).
15. We thank M. Palutan for the collaboration in this analysis.

**CP-VIOLATION PARAMETERS****Re( $\epsilon$ )**

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN
<b><math>1.596 \pm 0.013</math></b>	<sup>1</sup> AMBROSINO 06H KLOE	
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1.664 $\pm$ 0.010	<sup>2</sup> LAI 05A NA48	
<sup>1</sup> AMBROSINO 06H uses Bell-Steinberger relations with the following measurements: $B(K_L^0 \rightarrow \pi^+ \pi^-)$ in AMBROSINO 06F, $B(K_S^0 \rightarrow \pi^0 \pi^0 \pi^0)$ in AMBROSINO 05B, the $K_S^0$ -semileptonic charge asymmetry in AMBROSINO 06E, and $K^0$ -semileptonic results in ANGELOPOULOS 98F. <sup>2</sup> LAI 05A values are obtained through unitarity (Bell-Steinberger relations), improving determination of $\eta_{000}$ and combining other data from PDG 04 and APOSTOLAKIS 99B.		

**CPT-VIOLATION PARAMETERS**

In  $K^0$ - $\bar{K}^0$  mixing, if CP-violating interactions include a  $T$  conserving part then

$$|K_S\rangle = [|K_1\rangle + (\epsilon + \delta)|K_2\rangle] / \sqrt{1 + |\epsilon + \delta|^2}$$

$$|K_L\rangle = [|K_2\rangle + (\epsilon - \delta)|K_1\rangle] / \sqrt{1 + |\epsilon - \delta|^2}$$

where

$$|K_1\rangle = [|K^0\rangle + |\bar{K}^0\rangle] / \sqrt{2}$$

$$|K_2\rangle = [|K^0\rangle - |\bar{K}^0\rangle] / \sqrt{2}$$

and

$$|\bar{K}^0\rangle = CP|K^0\rangle.$$

The parameter  $\delta$  specifies the CPT-violating part.

Estimates of  $\delta$  are given below assuming the validity of the  $\Delta S = \Delta Q$  rule. See also THOMSON 95 for a test of CPT-symmetry conservation in  $K^0$  decays using the Bell-Steinberger relation.

**REAL PART OF  $\delta$** 

A nonzero value violates CPT invariance.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.51 \pm 2.25</math></b>		<sup>1</sup> ABOUZAID 11 KTEV		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.3 $\pm$ 2.7		<sup>2</sup> AMBROSINO 06H KLOE		
2.4 $\pm$ 2.8		<sup>3</sup> APOSTOLA... 99B RVUE		
2.9 $\pm$ 2.6 $\pm$ 0.6	1.3M	<sup>4</sup> ANGELOPO... 98F CPLR		
180 $\pm$ 200	6481	<sup>5</sup> DEMIDOV 95	$K_{\ell 3}$ reanalysis	

<sup>1</sup> ABOUZAID 11 uses Bell-Steinberger relations.

<sup>2</sup> AMBROSINO 06H uses Bell-Steinberger relations with the following measurements:  $B(K_L^0 \rightarrow \pi^+ \pi^-)$  in AMBROSINO 06F,  $B(K_S^0 \rightarrow \pi^0 \pi^0 \pi^0)$  in AMBROSINO 05B, the  $K_S^0$ -semileptonic charge asymmetry in AMBROSINO 06E, and  $K^0$ -semileptonic results in ANGELOPOULOS 98F.

<sup>3</sup> APOSTOLAKIS 99B assumes only unitarity and combines CPLEAR and other results.

<sup>4</sup> ANGELOPOULOS 98F use  $\Delta S = \Delta Q$ . If  $\Delta S = \Delta Q$  is not assumed, they find  $\text{Re}(\delta) = (3.0 \pm 3.3 \pm 0.6) \times 10^{-4}$ .

<sup>5</sup> DEMIDOV 95 reanalyzes data from HART 73 and NIEBERGALL 74.

**IMAGINARY PART OF  $\delta$** 

A nonzero value violates CPT invariance.

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-1.5 \pm 1.6</math></b>		<sup>1</sup> ABOUZAID 11 KTEV		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.4 $\pm$ 2.1		<sup>2</sup> AMBROSINO 06H KLOE		
-0.2 $\pm$ 2.0		<sup>3</sup> LAI 05A NA48		
2.4 $\pm$ 5.0		<sup>4</sup> APOSTOLA... 99B RVUE		
-90 $\pm$ 290 $\pm$ 100	1.3M	<sup>5</sup> ANGELOPO... 98F CPLR		
2100 $\pm$ 3700	6481	<sup>6</sup> DEMIDOV 95	$K_{\ell 3}$ reanalysis	

<sup>1</sup> ABOUZAID 11 uses Bell-Steinberger relations.

<sup>2</sup> AMBROSINO 06H uses Bell-Steinberger relations with the following measurements:  $B(K_L^0 \rightarrow \pi^+ \pi^-)$  in AMBROSINO 06F,  $B(K_S^0 \rightarrow \pi^0 \pi^0 \pi^0)$  in AMBROSINO 05B, the  $K_S^0$ -semileptonic charge asymmetry in AMBROSINO 06E, and  $K^0$ -semileptonic results in ANGELOPOULOS 98F.

<sup>3</sup> LAI 05A values are obtained through unitarity (Bell-Steinberger relations), improving determination of  $\eta_{000}$  and combining other data from PDG 04 and APOSTOLAKIS 99B.

<sup>4</sup> APOSTOLAKIS 99B assumes only unitarity and combines CPLEAR and other results.

<sup>5</sup> If  $\Delta S = \Delta Q$  is not assumed, ANGELOPOULOS 98F finds  $\text{Im}(\delta) = (-15 \pm 23 \pm 3) \times 10^{-3}$ .

<sup>6</sup> DEMIDOV 95 reanalyzes data from HART 73 and NIEBERGALL 74.

**Re( $y$ )**

A non-zero value would violate CPT invariance in  $\Delta S = \Delta Q$  amplitude. Re( $y$ ) is the following combination of  $K_{e3}$  decay amplitudes:

$$\text{Re}(y) = \text{Re} \left( \frac{A(K^0 \rightarrow e^- \pi^+ \bar{\nu}_e)^* - A(K^0 \rightarrow e^+ \pi^- \nu_e)}{A(K^0 \rightarrow e^- \pi^+ \bar{\nu}_e)^* + A(K^0 \rightarrow e^+ \pi^- \nu_e)} \right)$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN
<b><math>0.4 \pm 2.5</math></b>	13k	<sup>1</sup> AMBROSINO 06E KLOE	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.3 $\pm$ 3.1		<sup>2</sup> APOSTOLA... 99B CPLR	

<sup>1</sup> They use the PDG 04 for the  $K_L^0$  semileptonic charge asymmetry and PDG 04 (CP review, CPT NOT ASSUMED) for Re( $\epsilon$ ).

<sup>2</sup> Constrained by Bell-Steinberger (or unitarity) relation.

**Re( $x_-$ )**

A non-zero value would violate CPT invariance in decay amplitudes with  $\Delta S \neq \Delta Q$ .  $x_-$ , used here to define Re( $x_-$ ), and  $x_+$ , used below in the  $\Delta S = \Delta Q$  section are the following combinations of  $K_{e3}$  decay amplitudes:

$$x_{\pm} = \frac{1}{2} \left( \frac{A(\bar{K}^0 \rightarrow \pi^- e^+ \nu_e)}{A(K^0 \rightarrow \pi^- e^+ \nu_e)} \pm \frac{A(K^0 \rightarrow \pi^+ e^- \bar{\nu}_e)}{A(\bar{K}^0 \rightarrow \pi^+ e^- \bar{\nu}_e)} \right).$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-2.9 \pm 2.0</math></b>		<sup>1</sup> AMBROSINO 06H KLOE		
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.8 $\pm$ 2.5	13k	<sup>2</sup> AMBROSINO 06E KLOE		
-0.5 $\pm$ 3.0		<sup>3</sup> APOSTOLA... 99B CPLR	Strangeness tagged	
2 $\pm$ 13 $\pm$ 3	650k	ANGELOPO... 98F CPLR	Strangeness tagged	
<sup>1</sup> AMBROSINO 06H uses Bell-Steinberger relations with the following measurements: $B(K_L^0 \rightarrow \pi^+ \pi^-)$ in AMBROSINO 06F, $B(K_S^0 \rightarrow \pi^0 \pi^0 \pi^0)$ in AMBROSINO 05B, the $K_S^0$ -semileptonic charge asymmetry in AMBROSINO 06E, and $K^0$ -semileptonic results in ANGELOPOULOS 98F. <sup>2</sup> Uses PDG 04 for the $K_L^0$ semileptonic charge asymmetry and Re( $\delta$ ) from CPLEAR, ANGELOPOULOS 98F. <sup>3</sup> Constrained by Bell-Steinberger (or unitarity) relation.				

$$|m_{K^0} - m_{\bar{K}^0}| / m_{\text{average}}$$

A test of CPT invariance. "Our Evaluation" is described in the "Tests of Conservation Laws" section. It assumes CPT invariance in the decay and neglects some contributions from decay channels other than  $\pi\pi$ .

VALUE	CL%	DOCUMENT ID	TECN
<b><math>&lt; 6 \times 10^{-19}</math></b>	90	PDG 12	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$(-3 \pm 4) \times 10^{-18}$		<sup>1</sup> ANGELOPO... 99B RVUE	

<sup>1</sup> ANGELOPOULOS 99B assumes only unitarity and combines CPLEAR and other results.

$$(\Gamma_{K^0} - \Gamma_{\bar{K}^0}) / m_{\text{average}}$$

A test of CPT invariance.

VALUE	DOCUMENT ID	TECN
<b><math>(7.8 \pm 8.4) \times 10^{-18}</math></b>	<sup>1</sup> ANGELOPO... 99B RVUE	

<sup>1</sup> ANGELOPOULOS 99B assumes only unitarity and combines CPLEAR with other results. Correlated with  $(m_{K^0} - m_{\bar{K}^0}) / m_{\text{average}}$  with a correlation coefficient of  $-0.95$ .

**TESTS OF  $\Delta S = \Delta Q$  RULE****Re( $x_+$ )**

A non-zero value would violate the  $\Delta S = \Delta Q$  rule in CPT conserving transitions.  $x_+$  is defined above in the Re( $x_-$ ) section.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN
<b><math>-0.9 \pm 3.0</math> OUR AVERAGE</b>			
-2 $\pm$ 10		<sup>1</sup> BATLEY 07D NA48	
-0.5 $\pm$ 3.6	13k	<sup>2</sup> AMBROSINO 06E KLOE	
-1.8 $\pm$ 6.1		<sup>3</sup> ANGELOPO... 98D CPLR	

<sup>1</sup> Result obtained from the measurement  $\Gamma(K_S^0 \rightarrow \pi \nu) / \Gamma(K_L^0 \rightarrow \pi \nu) = 0.993 \pm 0.34$ , neglecting possible CPT non-invariance and using PDG 06 values of  $B(K_L^0 \rightarrow \pi \nu) = 0.4053 \pm 0.0015$ ,  $\tau_L = (5.114 \pm 0.021) \times 10^{-8}$  s and  $\tau_S = (0.8958 \pm 0.0005) \times 10^{-10}$  s.

<sup>2</sup> Re( $x_+$ ) can be shown to be equal to the following combination of rates:

$$\text{Re}(x_+) = \frac{1}{2} \frac{\Gamma(K_S^0 \rightarrow \pi \nu) - \Gamma(K_L^0 \rightarrow \pi \nu)}{\Gamma(K_S^0 \rightarrow \pi \nu) + \Gamma(K_L^0 \rightarrow \pi \nu)}$$

which is valid up to first order in terms violating CPT and/or the  $\Delta S = \Delta Q$  rule.

<sup>3</sup> Obtained neglecting CPT violating amplitudes.

# Meson Particle Listings

## $K^0, K_S^0$

### $K^0$ REFERENCES

TOMARADZE	14	PR D89 031501	A. Tomaradze <i>et al.</i>	(NWES, WAYN)
PDG	12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
ABOUZAID	11	PR D83 092001	E. Abouzaid <i>et al.</i>	(FNAL KTeV Collab.)
AMBROSINO	07B	JHEP 0712 073	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
BATLEY	07D	PL B653 145	J.R. Batley <i>et al.</i>	(CERN NA48 Collab.)
ABOUZAID	06	PRL 96 101801	E. Abouzaid <i>et al.</i>	(KTeV Collab.)
AMBROSINO	06E	PL B636 173	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AMBROSINO	06F	PL B638 140	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
AMBROSINO	06H	JHEP 0612 011	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
AMBROSINO	05B	PL B619 61	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
LAI	05A	PL B610 165	A. Lai <i>et al.</i>	(CERN NA48 Collab.)
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
LAI	03C	EPJ C30 33	A. Lai <i>et al.</i>	(CERN NA48 Collab.)
LAI	02	PL B533 196	A. Lai <i>et al.</i>	(CERN NA48 Collab.)
ANGELOPO... 01B	EPJ C22 55	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
ANGELOPO... 99B	PL B471 332	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
ANGELOPO... 99B	PL B456 297	A. Apostolakis <i>et al.</i>	(CLEAR Collab.)	
ANGELOPO... 98D	PL B444 38	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
Also	EPJ C22 55	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
ANGELOPO... 98E	PL B444 43	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
ANGELOPO... 98F	PL B444 52	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
Also	EPJ C22 55	A. Angelopoulos <i>et al.</i>	(CLEAR Collab.)	
DEMIDOV	95	PAN 58 968	V. Demidov, K. Gusev, E. Shabalín	(ITEP)
From YAF	58	1041.		
THOMSON	95	PR D51 1412	G.B. Thomson, Y. Zou	(RUTG)
BARKOV	87B	SJNP 46 630	L.M. Barkov <i>et al.</i>	(NOVO)
		Translated from YAF 46 1088.		
BARKOV	85B	JETPL 42 138	L.M. Barkov <i>et al.</i>	(NOVO)
		Translated from ZETFP 42 113.		
BLATNIK	79	LNC 24 39	S. Blatnik, J. Stahov, C.B. Lang	(TUZL, GRAZ)
MOLZON	78	PRL 41 1213	W.R. Molzon <i>et al.</i>	(EFI+)
NIEBERGALL	74	PL 49B 103	F. Niebergall <i>et al.</i>	(CERN, ORSAY, VIEN)
HART	73	NP B66 317	J.C. Hart <i>et al.</i>	(CAVE, RHEL)
FOETH	69B	PL 30B 276	H. Foeth <i>et al.</i>	(AACH, CERN, TORI)
HILL	68B	PR 168 1534	D.G. Hill <i>et al.</i>	(BNL, CMU)
FITCH	67	PR 164 1711	V.L. Fitch <i>et al.</i>	(PRIN)
BALTAY	66	PR 142 932	C. Baltay <i>et al.</i>	(YALE, BNL)
BURNSTEIN	65	PR 138 B895	R.A. Burnstein, H.A. Rubin	(UMD)
KIM	65B	PR 140B 1334	J.K. Kim, L. Kirsch, D. Miller	(COLU)
CHRISTENS...	64	PRL 13 138	J.H. Christenson <i>et al.</i>	(PRIN)
CRAWFORD	59	PRL 2 112	F.S. Crawford <i>et al.</i>	(LRL)
ROSENFELD	59	PRL 2 110	A.H. Rosenfeld, F.T. Solmitz, R.D. Tripp	(LRL)



$$I(J^P) = \frac{1}{2}(0^-)$$

### $K_S^0$ MEAN LIFE

For earlier measurements, beginning with BOLDT 58B, see our 1986 edition, Physics Letters **170B** 130 (1986).

OUR FIT is described in the note on “ $CP$  violation in  $K_L^0$  decays” in the  $K^0$  Particle Listings. The result labeled “OUR FIT Assuming  $CPT$ ” [“OUR FIT Not assuming  $CPT$ ”] includes all measurements except those with the comment “Not assuming  $CPT$ ” [“Assuming  $CPT$ ”]. Measurements with neither comment do not assume  $CPT$  and enter both fits.

VALUE ( $10^{-10}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.8954 ± 0.0004</b>	<b>OUR FIT</b>	Error includes scale factor of 1.1. Assuming $CPT$		
<b>0.89564 ± 0.00033</b>	<b>OUR FIT</b>	Not assuming $CPT$		
0.89589 ± 0.00070		1,2 ABOUZAID	11	KTEV Not assuming $CPT$
0.89623 ± 0.00047		1,3 ABOUZAID	11	KTEV Assuming $CPT$
0.89562 ± 0.00029 ± 0.00043	20M	4 AMBROSINO	11	KLOE Not assuming $CPT$
0.89598 ± 0.00048 ± 0.00051	16M	LAI	02C	NA48
0.8971 ± 0.0021		BERTANZA	97	NA31
0.8941 ± 0.0014 ± 0.0009		SCHWINGEN...	95	E773 Assuming $CPT$
0.8929 ± 0.0016		GIBBONS	93	E731 Assuming $CPT$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.8965 ± 0.0007		5 ALAVI-HARATI	03	KTEV Assuming $CPT$
0.8958 ± 0.0013		6 ALAVI-HARATI	03	KTEV Not assuming $CPT$
0.8920 ± 0.0044	214k	GROSSMAN	87	SPEC
0.905 ± 0.007		7 ARONSON	82B	SPEC
0.881 ± 0.009	26k	ARONSON	76	SPEC
0.8926 ± 0.0032 ± 0.0002		8 CARITHERS	75	SPEC
0.8937 ± 0.0048	6M	GEWENIGER	74B	ASPK
0.8958 ± 0.0045	50k	9 SKJEGGEST...	72	HBC
0.856 ± 0.008	19994	10 DONALD	68B	HBC
0.872 ± 0.009	20000	9,10 HILL	68	DBC

<sup>1</sup> The two ABOUZAID 11 values use the same full KTeV dataset from 1996, 1997, and 1999. The first enters the “assuming  $CPT$ ” fit and the second enters the “not assuming  $CPT$ ” fit.

<sup>2</sup> ABOUZAID 11 fit has  $\Delta m$ ,  $\tau_S$ ,  $\phi_e$ ,  $\text{Re}(\epsilon'/\epsilon)$ , and  $\text{Im}(\epsilon'/\epsilon)$  as free parameters. See  $\text{Im}(\epsilon'/\epsilon)$  in the “ $K_L^0$   $CP$  violation” section for correlation information.

<sup>3</sup> ABOUZAID 11 fit has  $\Delta m$  and  $\tau_S$  free but constrains  $\phi_e$  to the Superweak value, i.e. assumes  $CPT$ . This  $\tau_S$  value is correlated with their  $\Delta m = m_{K_L^0} - m_{K_S^0}$  measurement in the  $K_L^0$  listings. The correlation coefficient  $\rho(\tau_S, \Delta m) = -0.670$ .

<sup>4</sup> Fit to the proper time distribution.

<sup>5</sup> This ALAVI-HARATI 03 fit has  $\Delta m$  and  $\tau_S$  free but constrains  $\phi_{+-}$  to the Superweak value, i.e. assumes  $CPT$ . This  $\tau_S$  value is correlated with their  $\Delta m = m_{K_L^0} - m_{K_S^0}$  measurement in the  $K_L^0$  listings. The correlation coefficient  $\rho(\tau_S, \Delta m) = -0.396$ . Superseded by ABOUZAID 11.

<sup>6</sup> This ALAVI-HARATI 03 fit has  $\Delta m$ ,  $\phi_{+-}$ , and  $\tau_{K_S}$  free. See  $\phi_{+-}$  in the “ $K_L$   $CP$  violation” section for correlation information. Superseded by ABOUZAID 11.

<sup>7</sup> ARONSON 82 find that  $K_S^0$  mean life may depend on the kaon energy.

<sup>8</sup> CARITHERS 75 measures the  $\Delta m$  dependence of the total decay rate (inverse mean life) to be  $\Gamma(K_S^0) = [(1.122 \pm 0.004) + 0.16(\Delta m - 0.5348)/\Delta m]10^{10}/\text{s}$ , or, in terms of mean life, CARITHERS 75 measures  $\tau_S = (0.8913 \pm 0.0032) - 0.238 [\Delta m - 0.5348] (10^{-10} \text{ s})$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \hbar \text{ s}^{-1}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>9</sup> HILL 68 has been changed by the authors from the published value ( $0.865 \pm 0.009$ ) because of a correction in the shift due to  $\eta_{+-}$ . SKJEGGESTAD 72 and HILL 68 give detailed discussions of systematics encountered in this type of experiment.

<sup>10</sup> Pre-1971 experiments are excluded from the average because of disagreement with later more precise experiments.

### $K_S^0$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Hadronic modes</b>		
$\Gamma_1$ $\pi^0 \pi^0$	(30.69 ± 0.05) %	
$\Gamma_2$ $\pi^+ \pi^-$	(69.20 ± 0.05) %	
$\Gamma_3$ $\pi^+ \pi^- \pi^0$	( 3.5 $^{+1.1}_{-0.9}$ ) × 10 <sup>-7</sup>	
<b>Modes with photons or <math>\ell\bar{\ell}</math> pairs</b>		
$\Gamma_4$ $\pi^+ \pi^- \gamma$	[a,b] ( 1.79 ± 0.05) × 10 <sup>-3</sup>	
$\Gamma_5$ $\pi^+ \pi^- e^+ e^-$	( 4.79 ± 0.15) × 10 <sup>-5</sup>	
$\Gamma_6$ $\pi^0 \gamma \gamma$	[a] ( 4.9 ± 1.8 ) × 10 <sup>-8</sup>	
$\Gamma_7$ $\gamma \gamma$	( 2.63 ± 0.17) × 10 <sup>-6</sup>	S=3.0
<b>Semileptonic modes</b>		
$\Gamma_8$ $\pi^\pm e^\mp \nu_e$	[c] ( 7.04 ± 0.08) × 10 <sup>-4</sup>	
$\Gamma_9$ $\pi^\pm \mu^\mp \nu_\mu$	[c,d] ( 4.69 ± 0.05) × 10 <sup>-4</sup>	

### $CP$ violating ( $CP$ ) and $\Delta S = 1$ weak neutral current ( $S1$ ) modes

$\Gamma_{10}$ $3\pi^0$	$CP$	< 2.6	× 10 <sup>-8</sup>	CL=90%
$\Gamma_{11}$ $\mu^+ \mu^-$	$S1$	< 9	× 10 <sup>-9</sup>	CL=90%
$\Gamma_{12}$ $e^+ e^-$	$S1$	< 9	× 10 <sup>-9</sup>	CL=90%
$\Gamma_{13}$ $\pi^0 e^+ e^-$	$S1$	[a] ( 3.0 $^{+1.5}_{-1.2}$ )	× 10 <sup>-9</sup>	
$\Gamma_{14}$ $\pi^0 \mu^+ \mu^-$	$S1$	( 2.9 $^{+1.5}_{-1.2}$ )	× 10 <sup>-9</sup>	

[a] See the Particle Listings below for the energy limits used in this measurement.

[b] Most of this radiative mode, the low-momentum  $\gamma$  part, is also included in the parent mode listed without  $\gamma$ 's.

[c] The value is for the sum of the charge states or particle/antiparticle states indicated.

[d] Not a measurement. Calculated as  $0.666 \cdot B(\pi^\pm e^\mp \nu_e)$ .

### CONSTRAINED FIT INFORMATION

An overall fit to 4 branching ratios uses 5 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 0.1$  for 2 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-100		
$x_8$	-6	3	
$x_9$	-6	3	100
	$x_1$	$x_2$	$x_8$

### $K_S^0$ DECAY RATES

$\Gamma(\pi^\pm e^\mp \nu_e)$				$\Gamma_8$
VALUE ( $10^6 \text{ s}^{-1}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.1 $\pm$ 1.6	75	<sup>1</sup> AKHMETSHIN 99	CMD2	Tagged $K_S^0$ using $\phi \rightarrow K_L^0 K_S^0$
7.50 $\pm$ 0.08		<sup>2</sup> PDG	98	
seen		BURGUN	72	HBC $K^+ p \rightarrow K^0 p \pi^+$
9.3 $\pm$ 2.5		AUBERT	65	HLBC $\Delta S = \Delta Q$ , $CP$ cons. not assumed
<sup>1</sup> AKHMETSHIN 99 is from a measured branching ratio $B(K_S^0 \rightarrow \pi e \nu_e) = (7.2 \pm 1.4) \times 10^{-4}$ and $\tau_{K_S^0} = (0.8934 \pm 0.0008) \times 10^{-10} \text{ s}$ . Not independent of measured branching ratio.				
<sup>2</sup> PDG 98 from $K_L^0$ measurements, assuming that $\Delta S = \Delta Q$ in $K^0$ decay so that $\Gamma(K_S^0 \rightarrow \pi^\pm e^\mp \nu_e) = \Gamma(K_L^0 \rightarrow \pi^\pm e^\mp \nu_e)$ .				

See key on page 601

## Meson Particle Listings

 $K_S^0$ 

$\Gamma(\pi^\pm \mu^\mp \nu_\mu)$	$\Gamma_9$
VALUE ( $10^6 \text{ s}^{-1}$ )	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
5.25 ± 0.07	<sup>1</sup> PDG 98
<sup>1</sup> PDG 98 from $K_L^0$ measurements, assuming that $\Delta S = \Delta Q$ in $K^0$ decay so that $\Gamma(K_S^0 \rightarrow \pi^\pm \mu^\mp \nu_\mu) = \Gamma(K_L^0 \rightarrow \pi^\pm \mu^\mp \nu_\mu)$ .	

 $K_S^0$  BRANCHING RATIOS

## Hadronic modes

$\Gamma(\pi^0 \pi^0)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE	EVTS
0.3069 ± 0.0005 OUR FIT	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.335 ± 0.014	1066 BROWN 63 HLBC
0.288 ± 0.021	198 CHRETIEN 63 HLBC
0.30 ± 0.035	BROWN 61 HLBC

$\Gamma(\pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE	EVTS
0.6920 ± 0.0005 OUR FIT	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.670 ± 0.010	3447 DOYLE 69 HBC $\pi^- p \rightarrow \Lambda K^0$

$\Gamma(\pi^+ \pi^-)/\Gamma(\pi^0 \pi^0)$	$\Gamma_2/\Gamma_1$
VALUE	EVTS
2.255 ± 0.005 OUR FIT	
2.2549 ± 0.0054	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
2.2555 ± 0.0012 ± 0.0054	<sup>1</sup> AMBROSINO 06c KLOE
2.236 ± 0.003 ± 0.015	766k <sup>2</sup> ALOISIO 02b KLOE
2.11 ± 0.09	1315 EVERHART 76 WIRE $\pi^- p \rightarrow \Lambda K^0$
2.169 ± 0.094	16k COWELL 74 OSPK $\pi^- p \rightarrow \Lambda K^0$
2.16 ± 0.08	4799 HILL 73 DBC $K^+ d \rightarrow K^0 p p$
2.22 ± 0.10	3068 <sup>3</sup> ALITTI 72 HBC $K^+ p \rightarrow \pi^+ p K^0$
2.22 ± 0.08	6380 MORSE 72b DBC $K^+ n \rightarrow K^0 p$
2.10 ± 0.11	701 <sup>4</sup> NAGY 72 HLBC $K^+ n \rightarrow K^0 p$
2.22 ± 0.095	6150 <sup>5</sup> BALTAY 71 HBC $K p \rightarrow K^0 \text{neutrals}$
2.282 ± 0.043	7944 <sup>6</sup> MOFFETT 70 OSPK $K^+ n \rightarrow K^0 p$
2.12 ± 0.17	267 <sup>4</sup> BOZOKI 69 HLBC
2.285 ± 0.055	3016 <sup>6</sup> GOBBI 69 OSPK $K^+ n \rightarrow K^0 p$
2.10 ± 0.06	3700 MORFIN 69 HLBC $K^+ n \rightarrow K^0 p$

<sup>1</sup> This result combines AMBROSINO 06c KLOE 2001-02 data with ALOISIO 02b KLOE 2000 data.  $K_S^0 \rightarrow \pi^+ \pi^-$  fully inclusive.

<sup>2</sup> Includes radiative decays  $\pi^+ \pi^- \gamma$ .

<sup>3</sup> The directly measured quantity is  $K_S^0 \rightarrow \pi^+ \pi^- / \text{all } K^0 = 0.345 \pm 0.005$ .

<sup>4</sup> NAGY 72 is a final result which includes BOZOKI 69.

<sup>5</sup> The directly measured quantity is  $K_S^0 \rightarrow \pi^+ \pi^- / \text{all } \bar{K}^0 = 0.345 \pm 0.005$ .

<sup>6</sup> MOFFETT 70 is a final result which includes GOBBI 69.

$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE (units $10^{-7}$ )	EVTS
3.5 ± 1.1 OUR AVERAGE	
4.7 + 2.2 - 1.7 - 1.5	<sup>1</sup> BATLEY 05 NA48
2.5 + 1.3 - 1.0 - 0.6	500k <sup>2</sup> ADLER 97b CPLR
4.8 + 2.2 - 1.6 ± 1.1	<sup>3</sup> ZOU 96 E621
• • • We do not use the following data for averages, fits, limits, etc. • • •	
4.1 + 2.5 - 1.9 - 0.6	<sup>4</sup> ADLER 96e CPLR Sup. by ADLER 97b
3.9 + 5.4 - 1.8 - 0.7	<sup>5</sup> THOMSON 94 E621 Sup. by ZOU 96

<sup>1</sup> BATLEY 05 is obtained by measuring the interference parameters in  $K_S, K_L \rightarrow \pi^+ \pi^- \pi^0$ .  $\text{Re}(\lambda) = 0.038 \pm 0.008 \pm 0.006$  and  $\text{Im}(\lambda) = -0.013 \pm 0.005 \pm 0.004$ ; the correlation coeff. between  $\text{Re}(\lambda)$  and  $\text{Im}(\lambda)$  is 0.66 (statistical only).

<sup>2</sup> ADLER 97b find the  $CP$ -conserving parameters  $\text{Re}(\lambda) = (28 \pm 7 \pm 3) \times 10^{-3}$ ,  $\text{Im}(\lambda) = (-10 \pm 8 \pm 2) \times 10^{-3}$ . They estimate  $B(K_S^0 \rightarrow \pi^+ \pi^- \pi^0)$  from  $\text{Re}(\lambda)$  and the  $K_L^0$  decay parameters. See also ANGELOPOULOS 98c.

<sup>3</sup> ZOU 96 is from the the measured quantities  $|\rho_{+-0}| = 0.039 \pm 0.009 \pm 0.005$  and  $\phi_p = (-9 \pm 18)^\circ$ .

<sup>4</sup> ADLER 96e is from the measured quantities  $\text{Re}(\lambda) = 0.036 \pm 0.010 \pm 0.002 \pm 0.003$  and  $\text{Im}(\lambda)$  consistent with zero. Note that the quantity  $\lambda$  is the same as  $\rho_{+-0}$  used in other footnotes.

<sup>5</sup> THOMSON 94 calculates this branching ratio from their measurements  $|\rho_{+-0}| = 0.035 \pm 0.019 \pm 0.011 \pm 0.004$  and  $\phi_p = (-59 \pm 48)^\circ$  where  $|\rho_{+-0}| e^{i\phi_p} = A(K_S^0 \rightarrow \pi^+ \pi^- \pi^0, I = 2)/A(K_L^0 \rightarrow \pi^+ \pi^- \pi^0)$ .

Modes with photons or  $\ell\bar{\ell}$  pairs

$\Gamma(\pi^+ \pi^- \gamma)/\Gamma(\pi^+ \pi^-)$	$\Gamma_4/\Gamma_2$
VALUE (units $10^{-3}$ )	EVTS
2.59 ± 0.08 OUR AVERAGE	
2.56 ± 0.09	1286 RAMBERG 93 E731 $p_\gamma > 50 \text{ MeV}/c$
2.68 ± 0.15	<sup>1</sup> TAUREG 76 SPEC $p_\gamma > 50 \text{ MeV}/c$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
7.10 ± 0.22	3723 RAMBERG 93 E731 $p_\gamma > 20 \text{ MeV}/c$
3.0 ± 0.6	29 <sup>2</sup> BOBISUT 74 HLBC $p_\gamma > 40 \text{ MeV}/c$
2.8 ± 0.6	<sup>3</sup> BURGUN 73 HBC $p_\gamma > 50 \text{ MeV}/c$

<sup>1</sup> TAUREG 76 find direct emission contribution  $< 0.06$ , CL = 90%.

<sup>2</sup> BOBISUT 74 not included in average because  $p_\gamma$  cut differs. Estimates direct emission contribution to be 0.5 or less, CL = 95%.

<sup>3</sup> BURGUN 73 estimates that direct emission contribution is  $0.3 \pm 0.6$ .

$\Gamma(\pi^+ \pi^- e^+ e^-)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
VALUE (units $10^{-5}$ )	EVTS
4.79 ± 0.15 OUR AVERAGE	
4.83 ± 0.11 ± 0.14	23k <sup>1</sup> BATLEY 11 NA48 2002 data
4.69 ± 0.30	676 <sup>2</sup> LAI 03c NA48 1998+1999 data
• • • We do not use the following data for averages, fits, limits, etc. • • •	
4.71 ± 0.23 ± 0.22	620 <sup>2,3</sup> LAI 03c NA48 1999 data
4.5 ± 0.7 ± 0.4	56 LAI 00b NA48 1998 data

<sup>1</sup> BATLEY 11 reports  $[\Gamma(K_S^0 \rightarrow \pi^+ \pi^- e^+ e^-)/\Gamma_{\text{total}}] / [B(K_L^0 \rightarrow \pi^+ \pi^- \pi^0)] / [B(\pi^0 \rightarrow e^+ e^- \gamma)] = (3.28 \pm 0.06 \pm 0.04) \times 10^{-2}$  which we multiply by our best values  $B(K_L^0 \rightarrow \pi^+ \pi^- \pi^0) = (12.54 \pm 0.05) \times 10^{-2}$ ,  $B(\pi^0 \rightarrow e^+ e^- \gamma) = (1.174 \pm 0.035) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values. Also a limit on the absolute value of the interference between bremsstrahlung and E1 transition is given :  $< 4 \times 10^{-7}$  at 90% C.L.

<sup>2</sup> Uses normalization  $\text{BR}(K_L \rightarrow \pi^+ \pi^- \pi^0) * \text{BR}(\pi^0 \rightarrow e^+ e^- \gamma) = (1.505 \pm 0.047) \times 10^{-3}$  from our 2000 Edition.

<sup>3</sup> Second error is  $0.16(\text{syst}) \pm 0.15(\text{norm})$  combined in quadrature.

$\Gamma(\pi^0 \gamma \gamma)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$
VALUE (units $10^{-8}$ )	CL %
4.9 ± 1.6 ± 0.9	17
• • • We do not use the following data for averages, fits, limits, etc. • • •	
< 33	90

<sup>1</sup> Spectrum also measured and found consistent with the one generated by a constant matrix element.

$\Gamma(\gamma \gamma)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
VALUE (units $10^{-6}$ )	CL %
2.63 ± 0.17 OUR AVERAGE	
2.26 ± 0.12 ± 0.06	711
2.713 ± 0.063 ± 0.005	7.5k
• • • We do not use the following data for averages, fits, limits, etc. • • •	
2.58 ± 0.36 ± 0.22	149
2.2 ± 1.1	16
2.4 ± 0.9	35
< 13	90
2.4 ± 1.2	19
< 133	90

<sup>1</sup> AMBROSINO 08c reports  $(2.26 \pm 0.12 \pm 0.06) \times 10^{-6}$  from a measurement of  $[\Gamma(K_S^0 \rightarrow \gamma \gamma)/\Gamma_{\text{total}}] \times [B(K_S^0 \rightarrow \pi^0 \pi^0)]$  assuming  $B(K_S^0 \rightarrow \pi^0 \pi^0) = (30.69 \pm 0.05) \times 10^{-2}$ .

<sup>2</sup> LAI 03 reports  $[\Gamma(K_S^0 \rightarrow \gamma \gamma)/\Gamma_{\text{total}}] / [B(K_S^0 \rightarrow \pi^0 \pi^0)] = (8.84 \pm 0.18 \pm 0.10) \times 10^{-6}$  which we multiply by our best value  $B(K_S^0 \rightarrow \pi^0 \pi^0) = (30.69 \pm 0.05) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> BARR 95b result is calculated using  $B(K_L \rightarrow \gamma \gamma) = (5.86 \pm 0.17) \times 10^{-4}$ .

<sup>4</sup> BARR 95b quotes this as the combined BARR 95b + BURKHARDT 87 result after rescaling BURKHARDT 87 to use same branching ratios and lifetimes as BARR 95b.

## Semileptonic modes

$\Gamma(\pi^\pm e^\mp \nu_e)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
7.04 ± 0.08 OUR FIT	
7.04 ± 0.08 OUR AVERAGE	
7.046 ± 0.18 ± 0.16	<sup>1</sup> BATLEY 07b NA48 $K^0(\bar{K}^0)(t) \rightarrow \pi e \nu$
6.91 ± 0.34 ± 0.15	624 <sup>2</sup> ALOISIO 02 KLOE Tagged $K_S^0$ using $\phi \rightarrow K_L^0 K_S^0$
• • • We use the following data for averages but not for fits. • • •	
7.05 ± 0.09	13k <sup>3</sup> AMBROSINO 06e KLOE Not fitted
• • • We do not use the following data for averages, fits, limits, etc. • • •	
7.2 ± 1.4	75 AKHMETSHIN 99 CMD2 Tagged $K_S^0$ using $\phi \rightarrow K_L^0 K_S^0$

# Meson Particle Listings

## $K_S^0$

- <sup>1</sup> Reconstructed from  $K^0(\overline{K}^0)(t) \rightarrow \pi e \nu$  distributions using PDG values of  $B(K_L^0 \rightarrow \pi e \nu) = 0.4053 \pm 0.0015$ ,  $\tau_L = (5.114 \pm 0.021) \times 10^{-8}$  s and  $\tau_S = (0.8958 \pm 0.0005) \times 10^{-10}$  s.
- <sup>2</sup> Uses the PDG 00 value for  $B(K_S^0 \rightarrow \pi^+ \pi^-)$ .
- <sup>3</sup> Obtained by imposing  $\Sigma_i B(K_S^0 \rightarrow i) = 1$ , where  $i$  runs over all the four branching ratios  $\pi^+ \pi^-$ ,  $\pi^0 \pi^0$ ,  $\pi e \nu$ , and  $\pi \mu \nu$ . Input value of  $B(K_S^0 \rightarrow \pi^+ \pi^-) / B(K_S^0 \rightarrow \pi^0 \pi^0)$  from AMBROSINO 06c is used. To derive  $\Gamma(K_S^0 \rightarrow \pi^+ \mu \nu) / \Gamma(K_S^0 \rightarrow \pi^+ e \nu)$ , lepton universality is assumed, radiative corrections from ANDRE 07 are used, and phase space integrals are taken from KTeV, ALEXOPOULOS 04A. This branching fraction enters our fit via their  $\Gamma(\pi^\pm e^\mp \nu_e) / \Gamma(\pi^+ \pi^-)$  branching ratio measurement.

## $\Gamma(\pi^\pm \mu^\mp \nu_\mu) / \Gamma_{\text{total}}$ $\Gamma_9 / \Gamma$

The PDG 06 value below has not been measured but is computed to be 0.666 times the  $K_S \rightarrow \pi^\pm e^\mp \nu_e$  branching fraction. It is included in the fit that constrains the four branching ratios  $\pi^+ \pi^-$ ,  $\pi^0 \pi^0$ ,  $\pi e \nu$ , and  $\pi \mu \nu$  to sum to 1. This treatment, used by AMBROSINO 06E, is preferable to our previous practice of constraining the  $\pi^+ \pi^-$  and  $\pi^0 \pi^0$  modes to sum to 1. The 0.666 factor is obtained from AMBROSINO 06E and assumes lepton universality, radiative corrections from ANDRE 07, and phase space integrals from KTeV, ALEXOPOULOS 04A.

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>4.69 ± 0.06 OUR FIT</b>				
<b>4.691 ± 0.001 ± 0.056</b>		<sup>1</sup> PDG	06	calculated from $\pi^\pm e^\mp \nu_e$

- <sup>1</sup> The PDG 06 value is computed to be  $B_{PDG06}(\pi \mu \nu) = 0.666 B_{FIT}(\pi e \nu)$ . The first error specifies the arbitrarily small error,  $0.001 \times 10^{-4}$ , on  $B_{PDG06}(\pi \mu \nu)$  for fixed  $B_{FIT}(\pi e \nu)$ . The second error is that due to the uncertainty in  $B_{FIT}(\pi e \nu)$ .

## $\Gamma(\pi^\pm e^\mp \nu_e) / \Gamma(\pi^+ \pi^-)$ $\Gamma_8 / \Gamma_2$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN
<b>10.18 ± 0.12 OUR FIT</b>			
<b>10.19 ± 0.11 ± 0.07</b>	13k	AMBROSINO 06E	KLOE

### CP violating (CP) and $\Delta S = 1$ weak neutral current (S1) modes

## $\Gamma(3\pi^0) / \Gamma_{\text{total}}$ $\Gamma_{10} / \Gamma$

Violates CP conservation.

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.26</b>	90	590M	<sup>1</sup> BABUSCI	13c	KLOE $\phi \rightarrow K_L^0 K_S^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 1.2	90	37.8M	AMBROSINO 05B	KLOE	
< 7.4	90	4.9M	<sup>2</sup> LAI	05A	NA48
<140	90	7M	ACHASOV 99D	SND	
<190	90	17300	<sup>3</sup> ANGELOPO...	98B	CPLR
<370	90		BARMIN 83	HLBC	

- <sup>1</sup> BABUSCI 13c uses  $1.7 \text{ fb}^{-1}$  of data of  $\phi \rightarrow K_L^0 K_S^0$  decays with  $K_L^0$  interaction in the calorimeter, collected from 2004 to 2005. No candidate events were found in the data with an expected background of  $0.04^{+0.15}_{-0.03}$  events. Upper limit is obtained by normalizing to  $K_S^0 \rightarrow 2\pi^0$  decays.
- <sup>2</sup> LAI 05A value is obtained from their bound on  $|\eta_{000}|$  (not assuming CP T) and  $B(K_L^0 \rightarrow 3\pi^0) = 0.211 \pm 0.003$ , and PDG 04 values for  $K_L^0$  and  $K_S^0$  lifetimes. If CPT is assumed then  $B(K_S^0 \rightarrow 3\pi^0)_{CPT} < 2.3 \times 10^{-7}$  at 90% CL
- <sup>3</sup> ANGELOPOULOS 98B is from  $\text{Im}(\eta_{000}) = -0.05 \pm 0.12 \pm 0.05$ , assuming  $\text{Re}(\eta_{000}) = \text{Re}(\epsilon) = 1.635 \times 10^{-3}$  and using the value  $B(K_L^0 \rightarrow \pi^0 \pi^0 \pi^0) = 0.2112 \pm 0.0027$ .

## $\Gamma(\mu^+ \mu^-) / \Gamma_{\text{total}}$ $\Gamma_{11} / \Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN
<b>&lt; 9</b>	90	<sup>1</sup> AAIJ	13G LHCb
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< $3.2 \times 10^2$	90	GJESDAL 73	ASPK
< $7 \times 10^3$	90	HYAMS 69B	OSPK

- <sup>1</sup> AAIJ 13G uses  $1.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7$  TeV. They obtained  $B(K_S^0 \rightarrow \mu^+ \mu^-) < 11 \times 10^{-9}$  at 95% C.L.

## $\Gamma(e^+ e^-) / \Gamma_{\text{total}}$ $\Gamma_{12} / \Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.09</b>	90	<sup>1</sup> AMBROSINO 09A	KLOE	$e^+ e^- \rightarrow \phi \rightarrow K_S^0 K_L^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.4	90	ANGELOPO... 97	CPLR	
< 28	90	BLICK 94	CNTR	Hyperon facility
<100	90	BARMIN 86	XEBC	

- <sup>1</sup> AMBROSINO 09A reports  $< 0.09 \times 10^{-7}$  from a measurement of  $[\Gamma(K_S^0 \rightarrow e^+ e^-) / \Gamma_{\text{total}}] / [B(K_S^0 \rightarrow \pi^+ \pi^-)]$  assuming  $B(K_S^0 \rightarrow \pi^+ \pi^-) = (69.20 \pm 0.05) \times 10^{-2}$ .

## $\Gamma(\pi^0 e^+ e^-) / \Gamma_{\text{total}}$ $\Gamma_{13} / \Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.0^{+1.5}_{-1.2} \pm 0.2</math></b>		7	<sup>1</sup> BATLEY	03 NA48	$m_{ee} > 0.165 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 140	90		LAI	01	NA48
< 1100	90	0	BARR	93B	NA31
<45000	90		GIBBONS	88	E731

- <sup>1</sup> BATLEY 03 extrapolate also to the full kinematical region using a constant form factor and a vector matrix element. The resulting branching ratio is  $(5.8^{+2.9}_{-2.4}) \times 10^{-9}$ .

## $\Gamma(\pi^0 \mu^+ \mu^-) / \Gamma_{\text{total}}$ $\Gamma_{14} / \Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.9^{+1.5}_{-1.2} \pm 0.2</math></b>	6	<sup>1</sup> BATLEY	04A NA48	$NA48/1 K_S^0$ beam

- <sup>1</sup> Background estimate is  $0.22^{+0.18}_{-0.11}$  events. Branching ratio assumes a vector matrix element and unit form factor.

### $K_S^0$ FORM FACTORS

For discussion, see note on  $K_{f3}$  form factors in the  $K^\pm$  section of the Particle Listings above. Because the semileptonic branching fraction is smaller in  $K_S^0$  than  $K_L^0$  by the ratio of the mean lives, the  $K_S^0$  semileptonic form factor has so far been measured only in the  $K_{e3}$  mode using the linear expansion  $f_+(t) = f_+(0) (1 + \lambda_+ t / m_\pi^2)$ , which gives the vector form factor  $f_+(t)$  relative to its value at  $t = 0$ .

### $\lambda_+$ (LINEAR ENERGY DEPENDENCE OF $f_+$ IN $K_{e3}^0$ DECAY)

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN
<b>3.39 ± 0.41</b>	15k	AMBROSINO 06E	KLOE

### CP VIOLATION IN $K_S \rightarrow 3\pi$

Written 1996 by T. Nakada (Paul Scherrer Institute) and L. Wolfenstein (Carnegie-Mellon University).

The possible final states for the decay  $K^0 \rightarrow \pi^+ \pi^- \pi^0$  have isospin  $I = 0, 1, 2$ , and 3. The  $I = 0$  and  $I = 2$  states have  $CP = +1$  and  $K_S$  can decay into them without violating CP symmetry, but they are expected to be strongly suppressed by centrifugal barrier effects. The  $I = 1$  and  $I = 3$  states, which have no centrifugal barrier, have  $CP = -1$  so that the  $K_S$  decay to these requires CP violation.

In order to see CP violation in  $K_S \rightarrow \pi^+ \pi^- \pi^0$ , it is necessary to observe the interference between  $K_S$  and  $K_L$  decay, which determines the amplitude ratio

$$\eta_{+-0} = \frac{A(K_S \rightarrow \pi^+ \pi^- \pi^0)}{A(K_L \rightarrow \pi^+ \pi^- \pi^0)}. \quad (1)$$

If  $\eta_{+-0}$  is obtained from an integration over the whole Dalitz plot, there is no contribution from the  $I = 0$  and  $I = 2$  final states and a nonzero value of  $\eta_{+-0}$  is entirely due to CP violation.

Only  $I = 1$  and  $I = 3$  states, which are  $CP = -1$ , are allowed for  $K^0 \rightarrow \pi^0 \pi^0 \pi^0$  decays and the decay of  $K_S$  into  $3\pi^0$  is an unambiguous sign of CP violation. Similarly to  $\eta_{+-0}$ ,  $\eta_{000}$  is defined as

$$\eta_{000} = \frac{A(K_S \rightarrow \pi^0 \pi^0 \pi^0)}{A(K_L \rightarrow \pi^0 \pi^0 \pi^0)}. \quad (2)$$

If one assumes that CPT invariance holds and that there are no transitions to  $I = 3$  (or to nonsymmetric  $I = 1$  states), it can be shown that

$$\begin{aligned} \eta_{+-0} &= \eta_{000} \\ &= \epsilon + i \frac{\text{Im } a_1}{\text{Re } a_1}. \end{aligned} \quad (3)$$



# Meson Particle Listings

 $K_S^0, K_L^0$ 

MOFFETT	70	BAPS 15 512	R. Moffett <i>et al.</i>	(ROCH)
BOZOKI	69	PL 30B 498	G. Bozoki <i>et al.</i>	(BUDA)
DOYLE	69	Thesis UCRL 18139	J.C. Doyle	(LRL)
GOBBI	69	PRL 22 682	B. Gobbi <i>et al.</i>	(ROCH)
HYAMS	69B	PL 29B 521	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
MORFIN	69	PRL 23 660	J.G. Morfin, D. Sinclair	(MICH)
DONALD	68B	PL 27B 58	R.A. Donald <i>et al.</i>	(LIVP, CERN, IPNP+)
HILL	68	PR 171 1418	D.G. Hill <i>et al.</i>	(BNL, CMU)
AUBERT	65	PL 17 59	B. Aubert <i>et al.</i>	(EPOL, ORSAY)
BROWN	63B	PR 130 769	J.L. Brown <i>et al.</i>	(LRL, MICH)
CHRETIEN	63	PR 131 2208	M. Chretien <i>et al.</i>	(BRAN, BROW, HARV+)
BROWN	61	NC 19 1155	J.L. Brown <i>et al.</i>	(MICH)
BOLDT	58B	PRL 1 150	E. Boldt, D.O. Caldwell, Y. Pai	(MIT)

## OTHER RELATED PAPERS

LITTENBERG	93	ARNPS 43 729	L.S. Littenberg, G. Valencia	(BNL, FNAL)
BATTISTON	92	PRPL 214 293	Rare and Radiative Kaon Decays	(PGIA, CERN, TRSTT)
TRILLING	65B	UCRL 16473	R. Battiston <i>et al.</i>	(LRL)
CRAWFORD	62	CERN Conf. 827	Updated from 1965 Argonne Conference, page 115.	(LRL)
FITCH	61	NC 22 1160	G.N. Trilling	(LRL)
GOOD	61	PR 124 1223	F.S. Crawford	(PRIN+)
BIRGE	60	Rochester Conf. 601	V.L. Fitch, P.A. Piroue, R.B. Perkins	(LRL)
MULLER	60	PRL 4 418	R.H. Good <i>et al.</i>	(LRL, WISC)
			R.W. Birge <i>et al.</i>	(LRL, BNL)
			F. Muller <i>et al.</i>	(LRL, BNL)



$$I(J^P) = \frac{1}{2}(0^-)$$

$$m_{K_L^0} - m_{K_S^0}$$

For earlier measurements, beginning with GOOD 61 and FITCH 61, see our 1986 edition, Physics Letters **170B** 132 (1986).

OUR FIT is described in the note on “*CP* violation in  $K_L^0$  decays” in the  $K_L^0$  Particle Listings. The result labeled “OUR FIT Assuming *CPT*” [“OUR FIT Not assuming *CPT*”] includes all measurements except those with the comment “Not assuming *CPT*” [“Assuming *CPT*”]. Measurements with neither comment do not assume *CPT* and enter both fits.

VALUE ( $10^{10} \text{ h s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.5293 ± 0.0009 OUR FIT</b>	Error includes scale factor of 1.3. Assuming <i>CPT</i>		
<b>0.5289 ± 0.0010 OUR FIT</b>	Not assuming <i>CPT</i>		
0.52797 ± 0.00195	<sup>1,2</sup> ABOUZAID	11	KTEV Not assuming <i>CPT</i>
0.52699 ± 0.00123	<sup>1,3</sup> ABOUZAID	11	KTEV Assuming <i>CPT</i>
0.5240 ± 0.0044 ± 0.0033	APOSTOLA...	99c	CPLR $K^0\text{-}\bar{K}^0$ to $\pi^+\pi^-$
0.5297 ± 0.0030 ± 0.0022	<sup>4</sup> SCHWINGEN...	95	E773 20–160 GeV <i>K</i> beams
0.5286 ± 0.0028	<sup>5</sup> GIBBONS	93	E731 Assuming <i>CPT</i>
0.5257 ± 0.0049 ± 0.0021	<sup>4</sup> GIBBONS	93c	E731 Not assuming <i>CPT</i>
0.5340 ± 0.00255 ± 0.0015	<sup>6</sup> GEWENIGER	74c	SPEC Gap method
0.5334 ± 0.0040 ± 0.0015	<sup>6,7</sup> GJESDAL	74	SPEC Assuming <i>CPT</i>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.5261 ± 0.0015	<sup>8</sup> ALAVI-HARATI 03	KTEV	Assuming <i>CPT</i>
0.5288 ± 0.0043	<sup>9</sup> ALAVI-HARATI 03	KTEV	Not assuming <i>CPT</i>
0.5343 ± 0.0063 ± 0.0025	<sup>10</sup> ANGELOPO...	01	CPLR
0.5295 ± 0.0020 ± 0.0003	<sup>11</sup> ANGELOPO...	98D	CPLR Assuming <i>CPT</i>
0.5307 ± 0.0013	<sup>12</sup> ADLER	96c	RVUE
0.5274 ± 0.0029 ± 0.0005	<sup>11</sup> ADLER	95	CPLR Sup. by ANGELOPOU-LOS 98D
0.482 ± 0.014	<sup>13</sup> ARONSON	82B	SPEC $E=30\text{--}110$ GeV
0.534 ± 0.007	<sup>14</sup> CARNEGIE	71	ASPK Gap method
0.542 ± 0.006	<sup>14</sup> ARONSON	70	ASPK Gap method
0.542 ± 0.006	CULLEN	70	CNTR

<sup>1</sup> The two ABOUZAID 11 values use the same data. The first enters the “assuming *CPT*” fit and the second enters the “not assuming *CPT*” fit.

<sup>2</sup> ABOUZAID 11 fit has  $\Delta m$ ,  $\tau_S$ ,  $\phi_e$ ,  $\text{Re}(\epsilon'/\epsilon)$ , and  $\text{Im}(\epsilon'/\epsilon)$  as free parameters. See  $\text{Im}(\epsilon'/\epsilon)$  in the “ $K_L^0$  *CP* violation” section for correlation information.

<sup>3</sup> ABOUZAID 11 fit has  $\Delta m$  and  $\tau_S$  free but constrains  $\phi_e$  to the Superweak value, i.e. assumes *CPT*. See “ $K_S^0$  Mean Life” section for correlation information.

<sup>4</sup> Fits  $\Delta m$  and  $\phi_{+-}$  simultaneously. GIBBONS 93c systematic error is from B. Winstein via private communication. 20–160 GeV *K* beams.

<sup>5</sup> GIBBONS 93 value assume  $\phi_{+-} = \phi_{00} = \phi_{\text{SW}} = (43.7 \pm 0.2)^\circ$ , i.e. assumes *CPT*. 20–160 GeV *K* beams.

<sup>6</sup> These two experiments have a common systematic error due to the uncertainty in the momentum scale, as pointed out in WAHL 89.

<sup>7</sup> GJESDAL 74 uses charge asymmetry in  $K_{L3}^0$  decays.

<sup>8</sup> ALAVI-HARATI 03 fit  $\Delta m$  and  $\tau_{K_S^0}$  simultaneously.  $\phi_{+-}$  is constrained to the Superweak value, i.e. *CPT* is assumed. See “ $K_S^0$  Mean Life” section for correlation information. Superseded by ABOUZAID 11.

<sup>9</sup> ALAVI-HARATI 03 fit  $\Delta m$ ,  $\phi_{+-}$ , and  $\tau_{K_S}$  simultaneously. See  $\phi_{+-}$  in the “ $K_L$  *CP* violation” section for correlation information. Superseded by ABOUZAID 11.

<sup>10</sup> ANGELOPOULOS 01 uses strong interactions strangeness tagging at two different times.

<sup>11</sup> Uses  $\bar{K}_{e3}^0$  and  $K_{e3}^0$  strangeness tagging at production and decay. Assumes *CPT* conservation on  $\Delta S = -\Delta Q$  transitions.

<sup>12</sup> ADLER 96C is the result of a fit which includes nearly the same data as entered into the “OUR FIT” value above.

<sup>13</sup> ARONSON 82 find that  $\Delta m$  may depend on the kaon energy.

<sup>14</sup> ARONSON 70 and CARNEGIE 71 use  $K_S^0$  mean life =  $(0.862 \pm 0.006) \times 10^{-10}$  s. We have not attempted to adjust these values for the subsequent change in the  $K_S^0$  mean life or in  $\eta_{+-}$ .

## $K_L^0$ MEAN LIFE

VALUE ( $10^{-8}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.116 ± 0.021 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>5.099 ± 0.021 OUR AVERAGE</b>				
5.072 ± 0.011 ± 0.035	13M	<sup>1</sup> AMBROSINO 06	KLOE	$\sum_i B_i = 1$
5.092 ± 0.017 ± 0.025	15M	AMBROSINO 05c	KLOE	
5.154 ± 0.044	0.4M	VOSBURGH 72	CNTR	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.15 ± 0.14		DEVLIN 67	CNTR	

<sup>1</sup> AMBROSINO 06 uses  $\phi \rightarrow K_L K_S$  with  $K_L$  tagged by  $K_S \rightarrow \pi^+\pi^-$ . The four major  $K_L$  BR’s are measured, the small remainder ( $\pi^+\pi^-, \pi^0\pi^0, \gamma\gamma$ ) is taken from PDG 04. This KLOE  $K_L$  lifetime is obtained by imposing  $\sum_i B_i = 1$ . The correlation matrix among the four measured  $K_L$  BR’s and this  $K_L$  lifetime is

	$K_{e3}$	$K_{\mu 3}$	$3\pi^0$	$\pi^+\pi^-\pi^0$	$\tau_{K_L}$
$K_{e3}$	1	−0.25	−0.56	−0.07	0.25
$K_{\mu 3}$		1	−0.43	−0.20	0.33
$3\pi^0$			1	−0.39	−0.21
$\pi^+\pi^-\pi^0$				1	−0.39
$\tau_{K_L}$					1

These correlations are taken into account in our fit. The average of this KLOE mean life measurement and the independent KLOE measurement in AMBROSINO 05c is  $(5.084 \pm 0.023) \times 10^{-8}$  s.

## $K_L^0$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Semileptonic modes</b>		
$\Gamma_1 \quad \pi^\pm e^\mp \nu_{e3}$ Called $K_{e3}^0$ .	[a] (40.55 ± 0.11 ) %	S=1.7
$\Gamma_2 \quad \pi^\pm \mu^\mp \nu_{\mu 3}$ Called $K_{\mu 3}^0$ .	[a] (27.04 ± 0.07 ) %	S=1.1
$\Gamma_3 \quad (\pi \mu \text{atom}) \nu$	( 1.05 ± 0.11 ) $\times 10^{-7}$	
$\Gamma_4 \quad \pi^0 \pi^\pm e^\mp \nu$	[a] ( 5.20 ± 0.11 ) $\times 10^{-5}$	
$\Gamma_5 \quad \pi^\pm e^\mp \nu e^+ e^-$	[a] ( 1.26 ± 0.04 ) $\times 10^{-5}$	
<b>Hadronic modes, including Charge conjugation × Parity Violating (CPV) modes</b>		
$\Gamma_6 \quad 3\pi^0$	(19.52 ± 0.12 ) %	S=1.6
$\Gamma_7 \quad \pi^+\pi^-\pi^0$	(12.54 ± 0.05 ) %	
$\Gamma_8 \quad \pi^+\pi^-\pi^0$	CPV [b] ( 1.967 ± 0.010 ) $\times 10^{-3}$	S=1.5
$\Gamma_9 \quad \pi^0\pi^0$	CPV ( 8.64 ± 0.06 ) $\times 10^{-4}$	S=1.8

## Semileptonic modes with photons

$\Gamma_{10} \quad \pi^\pm e^\mp \nu_e \gamma$	[a,c,d] ( 3.79 ± 0.06 ) $\times 10^{-3}$	
$\Gamma_{11} \quad \pi^\pm \mu^\mp \nu_\mu \gamma$	( 5.65 ± 0.23 ) $\times 10^{-4}$	

## Hadronic modes with photons or $\ell\bar{\ell}$ pairs

$\Gamma_{12} \quad \pi^0\pi^0\gamma$	< 2.43	$\times 10^{-7}$	CL=90%
$\Gamma_{13} \quad \pi^+\pi^-\gamma$	[c,d] ( 4.15 ± 0.15 ) $\times 10^{-5}$		S=2.8
$\Gamma_{14} \quad \pi^+\pi^-\gamma(\text{DE})$	( 2.84 ± 0.11 ) $\times 10^{-5}$		S=2.0
$\Gamma_{15} \quad \pi^0 2\gamma$	[c] ( 1.273 ± 0.033 ) $\times 10^{-6}$		
$\Gamma_{16} \quad \pi^0 \gamma e^+ e^-$	( 1.62 ± 0.17 ) $\times 10^{-8}$		

## Other modes with photons or $\ell\bar{\ell}$ pairs

$\Gamma_{17}$	$2\gamma$	$(5.47 \pm 0.04) \times 10^{-4}$	S=1.1
$\Gamma_{18}$	$3\gamma$	$< 7.4 \times 10^{-8}$	CL=90%
$\Gamma_{19}$	$e^+e^-\gamma$	$(9.4 \pm 0.4) \times 10^{-6}$	S=2.0
$\Gamma_{20}$	$\mu^+\mu^-\gamma$	$(3.59 \pm 0.11) \times 10^{-7}$	S=1.3
$\Gamma_{21}$	$e^+e^-\gamma\gamma$	$[c] (5.95 \pm 0.33) \times 10^{-7}$	
$\Gamma_{22}$	$\mu^+\mu^-\gamma\gamma$	$[c] (1.0 \pm 0.8) \times 10^{-8}$	

## Charge conjugation × Parity (CP) or Lepton Family number (LF) violating modes, or $\Delta S = 1$ weak neutral current (S1) modes

$\Gamma_{23} \quad \mu^+ \mu^-$	S1 ( 6.84 ± 0.11 ) $\times 10^{-9}$		
$\Gamma_{24} \quad e^+ e^-$	S1 ( 9 ± 6 ) $\times 10^{-12}$		
$\Gamma_{25} \quad \pi^+\pi^-\pi^+e^-$	S1 [c] ( 3.11 ± 0.19 ) $\times 10^{-7}$		
$\Gamma_{26} \quad \pi^0\pi^0 e^+ e^-$	S1 < 6.6	$\times 10^{-9}$	CL=90%
$\Gamma_{27} \quad \pi^0\pi^0 \mu^+ \mu^-$	S1 < 9.2	$\times 10^{-11}$	CL=90%
$\Gamma_{28} \quad \mu^+ \mu^- e^+ e^-$	S1 ( 2.69 ± 0.27 ) $\times 10^{-9}$		
$\Gamma_{29} \quad e^+ e^- e^+ e^-$	S1 ( 3.56 ± 0.21 ) $\times 10^{-8}$		



# Meson Particle Listings

$K_L^0$

$\Gamma(\pi^0\pi^\pm e^\mp\nu)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	
<b>5.20±0.11 OUR AVERAGE</b>					
5.21±0.07±0.09		5402	BATLEY	04	NA48
5.16±0.20±0.22		729	MAKOFF	93	E731
• • • We do not use the following data for averages, fits, limits, etc.					• • •
6.2 ±2.0		16	CARROLL	80c	SPEC
< 220		90	<sup>1</sup> DONALDSON	74	SPEC
<sup>1</sup> DONALDSON 74 uses $K_L^0 \rightarrow \pi^+\pi^-\pi^0$ /(all $K_L^0$ ) decays = 0.126.					

$\Gamma(\pi^\pm e^\mp\nu e^+e^-)/\Gamma(\pi^+\pi^-\pi^0)$					$\Gamma_5/\Gamma_7$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>10.02±0.17±0.29</b>	19k	<sup>1</sup> ABOUZAID	07c	KTEV $M_{ee}>5\text{ MeV}$ , $E_{ee}^*>30\text{ MeV}$	
<sup>1</sup> $E_{ee}^*$ is the energy of the $e^+e^-$ pair in the kaon rest frame. ABOUZAID 07c reports $[\Gamma(K_L^0 \rightarrow \pi^\pm e^\mp\nu e^+e^-)/\Gamma(K_L^0 \rightarrow \pi^+\pi^-\pi^0)] / [B(\pi^0 \rightarrow e^+e^-\gamma)] = (8.54 \pm 0.07 \pm 0.13) \times 10^{-3}$ which we multiply by our best value $B(\pi^0 \rightarrow e^+e^-\gamma) = (1.174 \pm 0.035) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

————— Hadronic modes, —————  
including Charge conjugation×Parity Violating (CPV) modes —————

$\Gamma(3\pi^0)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.1952±0.0012 OUR FIT</b>	Error includes scale factor of 1.6.				
<b>0.1969±0.0026 OUR AVERAGE</b>	Error includes scale factor of 2.0.				
• • • We use the following data for averages but not for fits.	• • •				
0.1997±0.0003±0.0019	13M	<sup>1</sup> AMBROSINO	06	KLOE	Not fitted
0.1945±0.0018		<sup>1</sup> ALEXOPOU...	04	KTEV	Not fitted
<sup>1</sup> We exclude these $B(K_L \rightarrow 3\pi^0)$ measurements from our fit because the authors have constrained $K_L$ branching fractions to sum to one. It enters our fit via the other measurements from the experiment and their correlations, along with our constraint that the fitted branching fractions sum to one.					

$\Gamma(3\pi^0)/\Gamma(\pi^\pm e^\mp\nu_e)$					$\Gamma_6/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.481 ±0.004 OUR FIT</b>	Error includes scale factor of 1.8.				
• • • We use the following data for averages but not for fits.	• • •				
<b>0.4782±0.0014±0.0053</b>	209K	<sup>1</sup> ALEXOPOU...	04	KTEV	Not in fit
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
0.545 ±0.004 ±0.009	38k	KREUTZ	95	NA31	
<sup>1</sup> This measurement enters the fit via their separate measurements of these two modes.					

$\Gamma(3\pi^0)/[\Gamma(\pi^\pm e^\mp\nu_e) + \Gamma(\pi^\pm\mu^\mp\nu_\mu) + \Gamma(\pi^+\pi^-\pi^0)]$					$\Gamma_6/(\Gamma_1+\Gamma_2+\Gamma_7)$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.2436±0.0018 OUR FIT</b>	Error includes scale factor of 1.6.				
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
0.251 ±0.014	549	BUDAGOV	68	HLBC	ORSAY measur.
0.277 ±0.021	444	BUDAGOV	68	HLBC	Ecole polytec.meas
0.31 ± <sup>+0.07</sup> <sub>-0.06</sub>	29	KULYUKINA	68	CC	
0.24 ±0.08	24	ANIKINA	64	CC	

$\Gamma(3\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$					$\Gamma_6/\Gamma_7$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.557±0.012 OUR FIT</b>	Error includes scale factor of 1.3.				
• • • We use the following data for averages but not for fits.	• • •				
<b>1.582±0.027</b>	13M	<sup>1</sup> AMBROSINO	06	KLOE	Not in fit
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
1.611±0.014±0.034	28k	KREUTZ	95	NA31	
1.65 ±0.07	883	BARMIN	72b	HLBC	Error statistical only
1.80 ±0.13	1010	BUDAGOV	68	HLBC	
2.0 ±0.6	188	ALEKSANYAN	64b	FBC	
<sup>1</sup> AMBROSINO 06 enters the fit via their separate measurements of these two modes.					

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.1254±0.0005 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>0.1255±0.0006 OUR AVERAGE</b>					
0.1263±0.0004±0.0011	13M	<sup>1</sup> AMBROSINO	06	KLOE	
0.1252±0.0007		<sup>2</sup> ALEXOPOU...	04	KTEV	
<sup>1</sup> There are correlations between these five KLOE measurements: $B(K_L \rightarrow \pi e\nu)$ , $B(K_L \rightarrow \pi\mu\nu)$ , $B(K_L \rightarrow 3\pi^0)$ , $B(K_L \rightarrow \pi^+\pi^-\pi^0)$ , and $\tau_{K_L}$ measured in AMBROSINO 06. See the footnote for the $\tau_{K_L}$ measurement for the correlation matrix.					
<sup>2</sup> For correlations with other ALEXOPOULOS 04 measurements, see the footnote with their $B(K_L \rightarrow \pi e\nu)$ measurement.					

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\pi^\pm e^\mp\nu_e)$					$\Gamma_7/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.3092±0.0016 OUR FIT</b>	Error includes scale factor of 1.1.				
• • • We use the following data for averages but not for fits.	• • •				
<b>0.3078±0.0005±0.0017</b>	799K	<sup>1</sup> ALEXOPOU...	04	KTEV	Not in fit
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
0.336 ±0.003 ±0.007	28k	KREUTZ	95	NA31	
<sup>1</sup> This measurement enters the fit via their separate measurements for the two modes.					

$\Gamma(\pi^+\pi^-\pi^0)/[\Gamma(\pi^\pm e^\mp\nu_e) + \Gamma(\pi^\pm\mu^\mp\nu_\mu) + \Gamma(\pi^+\pi^-\pi^0)]$					$\Gamma_7/(\Gamma_1+\Gamma_2+\Gamma_7)$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.1565±0.0006 OUR FIT</b>	Error includes scale factor of 1.1.				
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
0.163 ±0.003	6499	CHO	77	HBC	
0.1605±0.0038	1590	ALEXANDER	73b	HBC	
0.146 ±0.004	3200	BRANDENB...	73	HBC	
0.159 ±0.010	558	EVANS	73	HLBC	
0.167 ±0.016	1402	KULYUKINA	68	CC	
0.161 ±0.005		HOPKINS	67	HBC	
0.162 ±0.015	126	HAWKINS	66	HBC	
0.159 ±0.015	326	ASTBURY	65b	CC	
0.178 ±0.017	566	GUIDONI	65	HBC	
0.144 ±0.004	1729	HOPKINS	65	HBC	See HOPKINS 67

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$
Violates CP conservation.					
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.967±0.010 OUR FIT</b>	Error includes scale factor of 1.5.				
<b>1.975±0.012</b>		<sup>1</sup> ALEXOPOU...	04	KTEV	
<sup>1</sup> For correlations with other ALEXOPOULOS 04 measurements, see the footnote with their $B(K_L \rightarrow \pi e\nu)$ measurement.					

$\Gamma(\pi^+\pi^-)/\Gamma(\pi^\pm e^\mp\nu_e)$					$\Gamma_8/\Gamma_1$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.849±0.020 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>4.840±0.020 OUR AVERAGE</b>					
4.826±0.022±0.016	47k	<sup>1</sup> LAI	07	NA48	
• • • We use the following data for averages but not for fits.	• • •				
4.856±0.017±0.023	84k	<sup>2</sup> ALEXOPOU...	04	KTEV	Not in fit
<sup>1</sup> The LAI 07 central value of $4.835 \times 10^{-3}$ has been reduced by 0.19% to $4.826 \times 10^{-3}$ to subtract the contribution from the direct emission mode $K_L^0 \rightarrow \pi^+\pi^-\gamma$ (DE).					
<sup>2</sup> This measurement enters the fit via their separate measurements for the two modes.					

$[\Gamma(\pi^+\pi^-) + \Gamma(\pi^+\pi^-\gamma(\text{DE}))]/\Gamma(\pi^\pm\mu^\mp\nu_\mu)$					$(\Gamma_8+\Gamma_{14})/\Gamma_2$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>7.38 ±0.04 OUR FIT</b>	Error includes scale factor of 1.4.				
<b>7.275±0.042±0.054</b>	45k	<sup>1</sup> AMBROSINO	06f	KLOE	
<sup>1</sup> Fully inclusive. Taking $B(K_L^0 \rightarrow \pi\mu\nu)$ from KLOE, AMBROSINO 06, $B(K_L^0 \rightarrow \pi^+\pi^- + \pi^+\pi^-\gamma(\text{DE})) = (1.963 \pm 0.012 \pm 0.017) \times 10^{-3}$ is obtained.					

$\Gamma(\pi^+\pi^-)/[\Gamma(\pi^\pm e^\mp\nu_e) + \Gamma(\pi^\pm\mu^\mp\nu_\mu)]$					$\Gamma_8/(\Gamma_1+\Gamma_2)$
Violates CP conservation.					
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.909±0.013 OUR FIT</b>	Error includes scale factor of 1.3.				
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
3.13 ±0.14	1687	COUPAL	85	SPEC	$\eta_{+-}=-2.28 \pm 0.06$
3.04 ±0.14	2703	DEVOE	77	SPEC	$\eta_{+-}=-2.25 \pm 0.05$
2.51 ±0.23	309	<sup>1</sup> DEBOUARD	67	OSPK	$\eta_{+-}=-2.00 \pm 0.09$
2.35 ±0.19	525	<sup>1</sup> FITCH	67	OSPK	$\eta_{+-}=-1.94 \pm 0.08$
<sup>1</sup> Old experiments excluded from fit. See subsection on $\eta_{+-}$ in section on "PARAMETERS FOR $K_L^0 \rightarrow 2\pi$ DECAY" below for average $\eta_{+-}$ of these experiments and for note on discrepancy.					

$\Gamma(\pi^\pm e^\mp\nu_e)/\Gamma(2\text{ tracks})$					$\Gamma_1/(\Gamma_1+\Gamma_2+0.03508\Gamma_6+\Gamma_7+\Gamma_8)$
$\Gamma(2\text{ tracks}) = \Gamma(\pi^\pm e^\mp\nu_e) + \Gamma(\pi^\pm\mu^\mp\nu_\mu) + 0.03508\Gamma(3\pi^0) + \Gamma(\pi^+\pi^-\pi^0) + \Gamma(\pi^+\pi^-)$ where 0.03508 is the fraction of $3\pi^0$ events with one Dalitz decay ( $\pi^0 \rightarrow \gamma e^\pm e^-$ ).					

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.5006±0.0009 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>0.4978±0.0035</b>	6.8M	LAI	04b	NA48

$\Gamma(\pi^+\pi^-)/[\Gamma(\pi^\pm e^\mp\nu_e) + \Gamma(\pi^\pm\mu^\mp\nu_\mu) + \Gamma(\pi^+\pi^-\pi^0)]$					$\Gamma_8/(\Gamma_1+\Gamma_2+\Gamma_7)$
Violates CP conservation.					
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.454±0.011 OUR FIT</b>	Error includes scale factor of 1.3.				
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
2.60 ±0.07	4200	<sup>1</sup> MESSNER	73	ASPK	$\eta_{+-} = 2.23 \pm 0.05$
<sup>1</sup> From same data as $\Gamma(\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$ MESSNER 73, but with different normalization.					

$\Gamma(\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$					$\Gamma_8/\Gamma_7$
Violates CP conservation.					
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.568±0.010 OUR FIT</b>	Error includes scale factor of 1.3.				
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
1.64 ±0.04	4200	MESSNER	73	ASPK	$\eta_{+-} = 2.23$

$\Gamma(\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
Violates CP conservation.					
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.864±0.006 OUR FIT</b>	Error includes scale factor of 1.8.				
<b>0.865±0.012</b>		<sup>1</sup> ALEXOPOU...	04	KTEV	
<sup>1</sup> For correlations with other ALEXOPOULOS 04 measurements, see the footnote with their $B(K_L \rightarrow \pi e\nu)$ measurement.					



$\Gamma(\pi^0\pi^0)/\Gamma(\pi^+\pi^-)$ 

Violates CP conservation.

 $\Gamma_9/\Gamma_8$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.4395 ± 0.0023 OUR FIT</b>				Error includes scale factor of 2.0.
<b>0.4390 ± 0.0012</b>		ETAFIT	16	

 $\Gamma(\pi^0\pi^0)/\Gamma(3\pi^0)$ 

Violates CP conservation.

 $\Gamma_9/\Gamma_6$ 

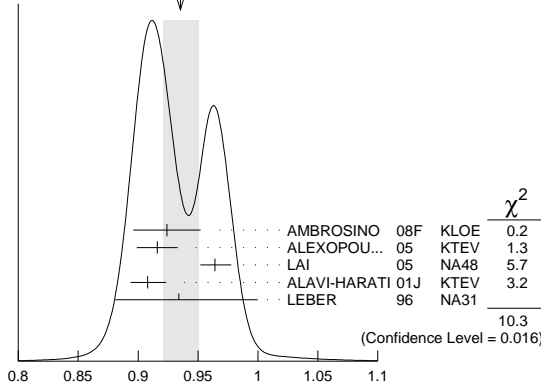
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.443 ± 0.004 OUR FIT</b>				Error includes scale factor of 2.1.
• • • We use the following data for averages but not for fits. • • •				
<b>0.4446 ± 0.0016 ± 0.0019</b>	100K	<sup>1</sup> ALEXOPOU...	04 KTEV	Not in fit
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.37 ± 0.08	29	BARMIN	70 HLBC	$\eta_{00}=2.02 \pm 0.23$
0.32 ± 0.15	30	BUDAGOV	70 HLBC	$\eta_{00}=1.9 \pm 0.5$
0.46 ± 0.11	57	BANNER	69 OSPK	$\eta_{00}=2.2 \pm 0.3$

<sup>1</sup> This measurement enters the fit via their separate measurements for the two modes.

## Semileptonic modes with photons

 $\Gamma(\pi^\pm e^\mp \nu_e \gamma)/\Gamma(\pi^\pm e^\mp \nu_e)$  $\Gamma_{10}/\Gamma_1$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.935 ± 0.015 OUR AVERAGE</b>				Error includes scale factor of 1.9. See the ideogram below.
0.924 ± 0.023 ± 0.016	9k	<sup>1</sup> AMBROSINO 08F	KLOE	$E_\gamma^* > 30$ MeV, $\theta_{e\gamma}^* > 20^\circ$
0.916 ± 0.017	4309	<sup>2</sup> ALEXOPOU... 05	KTEV	$E_\gamma^* > 30$ MeV, $\theta_{e\gamma}^* > 20^\circ$
0.964 ± 0.008 $^{+0.011}_{-0.009}$	19K	LAI 05	NA48	$E_\gamma^* > 30$ MeV, $\theta_{e\gamma}^* > 20^\circ$
0.908 ± 0.008 $^{+0.013}_{-0.012}$	15k	ALAVI-HARATI 01J	KTEV	$E_\gamma^* \geq 30$ MeV, $\theta_{e\gamma}^* \geq 20^\circ$
0.934 ± 0.036 $^{+0.055}_{-0.039}$	1384	LEBER 96	NA31	$E_\gamma^* \geq 30$ MeV, $\theta_{e\gamma}^* \geq 20^\circ$

<sup>1</sup> Direct emission contribution measured  $\langle X \rangle = -2.3 \pm 1.3 \pm 1.4$ .<sup>2</sup> Also measured cut  $E_\gamma^* > 10$  MeV,  $\theta_{e\gamma}^* > 0^\circ$  14221 evts:  $\Gamma(\pi^\pm e^\mp \nu_e \gamma) / \Gamma(\pi^\pm e^\mp \nu_e) = (4.942 \pm 0.062)\%$ .WEIGHTED AVERAGE  
0.935 ± 0.015 (Error scaled by 1.9) $\Gamma(\pi^\pm e^\mp \nu_e \gamma)/\Gamma(\pi^\pm e^\mp \nu_e)$  (units  $10^{-2}$ ) $\Gamma(\pi^\pm \mu^\mp \nu_\mu \gamma)/\Gamma(\pi^\pm \mu^\mp \nu_\mu)$  $\Gamma_{11}/\Gamma_2$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.09 ± 0.08 OUR AVERAGE</b>				
2.09 ± 0.09		<sup>1</sup> ALEXOPOU... 05	KTEV	$E_\gamma^* > 30$ MeV
2.08 ± 0.17 $^{+0.16}_{-0.21}$	252	BENDER 98	NA48	$E_\gamma^* \geq 30$ MeV

<sup>1</sup> Also measured cut  $E_\gamma^* > 10$  MeV, 1385 evts:  $\Gamma(\pi^\pm \mu^\mp \nu_\mu \gamma) / \Gamma(\pi^\pm \mu^\mp \nu_\mu) = (0.530 \pm 0.014 \pm 0.012)\%$ .Hadronic modes with photons or  $\ell\bar{\ell}$  pairs $\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$  $\Gamma_{12}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.243</b>	90	ABOUZAID 08B	KTEV	$K_L^0 \rightarrow \pi^0\pi_D^0\gamma, \pi_D^0 \rightarrow e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 5.6	90	BARR 94	NA31	
< 230	90	ROBERTS 94	E799	

 $\Gamma(\pi^+\pi^-\gamma)/\Gamma(\pi^+\pi^-\pi^0)$  $\Gamma_{13}/\Gamma_7$ 

For earlier limits see our 1992 edition Physical Review D45 S1 (1992).

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.23 ± 0.13	516	<sup>1,2</sup> CARROLL 80B	SPEC	$E_\gamma^* > 20$ MeV
2.33 ± 0.23	546	<sup>1,3</sup> CARROLL 80B	SPEC	
3.56 ± 0.26	1062	<sup>1,4</sup> CARROLL 80B	SPEC	$E_\gamma^* > 20$ MeV

<sup>1</sup> CARROLL 80B quotes  $B(\pi^+\pi^-\gamma)$  using normalization  $B(\pi^+\pi^-\pi^0) = 0.1239$ . We divide by this value to obtain their measured  $\Gamma(\pi^+\pi^-\gamma) / \Gamma(\pi^+\pi^-\pi^0)$ .<sup>2</sup> Internal Bremsstrahlung component only.<sup>3</sup> Direct  $\gamma$  emission component only.<sup>4</sup> Both IB and DE components. $\Gamma(\pi^+\pi^-\gamma)/\Gamma(\pi^+\pi^-)$  $\Gamma_{13}/\Gamma_8$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.11 ± 0.08 OUR FIT</b>				Error includes scale factor of 2.9.
<b>2.11 ± 0.08 OUR AVERAGE</b>				Error includes scale factor of 2.9.
2.08 ± 0.02 ± 0.02	8669	<sup>1</sup> ALAVI-HARATI 01B	KTEV	$E_\gamma^* > 20$ MeV
2.30 ± 0.07	3136	RAMBERG 93	E731	$E_\gamma^* > 20$ MeV

<sup>1</sup> ALAVI-HARATI 01B includes both Direct Emission (DE) and Inner Bremsstrahlung (IB) processes. $\Gamma(\pi^+\pi^-\gamma(\text{DE}))/\Gamma(\pi^+\pi^-\gamma)$  $\Gamma_{14}/\Gamma_{13}$ These values assume that  $\Gamma(K_L^0 \rightarrow \pi^+\pi^-\gamma) = \Gamma(K_L^0 \rightarrow \pi^+\pi^-\gamma(\text{DE})) + \Gamma(K_L^0 \rightarrow \pi^+\pi^-\gamma(\text{IB}))$ , the sum of widths for the direct emission (DE) and inner bremsstrahlung (IE) processes, with no IB-DE interference. DE assumes a form factor as described in RAMBERG 93.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.684 ± 0.009 OUR FIT</b>				
<b>0.684 ± 0.009 OUR AVERAGE</b>				
0.689 ± 0.021	111k	ABOUZAID 06A	KTEV	$E_\gamma^* > 20$ MeV
0.683 ± 0.011	8669	ALAVI-HARATI 01B	KTEV	$E_\gamma^* > 20$ MeV
0.685 ± 0.041	3136	RAMBERG 93	E731	$E_\gamma^* > 20$ MeV

 $\Gamma(\pi^0 2\gamma)/\Gamma_{\text{total}}$  $\Gamma_{15}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.273 ± 0.033 OUR AVERAGE</b>					
1.28 ± 0.06 ± 0.01	1.4k		<sup>1</sup> ABOUZAID 08	KTEV	
1.27 ± 0.04 ± 0.01	2.5k		<sup>2</sup> LAI 02B	NA48	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.68 ± 0.07 ± 0.08	884		<sup>3</sup> ALAVI-HARATI 99B	KTEV	
1.7 ± 0.2 ± 0.2	63		<sup>4</sup> BARR 92	NA31	
1.86 ± 0.60 ± 0.60	60		PAPADIMITR...91	E731	$m_{\gamma\gamma} > 280$ MeV
< 5.1	90		PAPADIMITR...91	E731	$m_{\gamma\gamma} < 264$ MeV
2.1 ± 0.6	14		<sup>5</sup> BARR 90c	NA31	$m_{\gamma\gamma} > 280$ MeV

<sup>1</sup> ABOUZAID 08 reports  $(1.29 \pm 0.03 \pm 0.05) \times 10^{-6}$  from a measurement of  $[\Gamma(K_L^0 \rightarrow \pi^0 2\gamma)/\Gamma_{\text{total}}] / [B(K_L^0 \rightarrow \pi^0\pi^0)]$  assuming  $B(K_L^0 \rightarrow \pi^0\pi^0) = (8.69 \pm 0.04) \times 10^{-4}$ , which we rescale to our best value  $B(K_L^0 \rightarrow \pi^0\pi^0) = (8.64 \pm 0.06) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>2</sup> LAI 02B reports  $[\Gamma(K_L^0 \rightarrow \pi^0 2\gamma)/\Gamma_{\text{total}}] / [B(K_L^0 \rightarrow \pi^0\pi^0)] = (1.467 \pm 0.032 \pm 0.032) \times 10^{-3}$  which we multiply by our best value  $B(K_L^0 \rightarrow \pi^0\pi^0) = (8.64 \pm 0.06) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. They also find that  $B(\pi^0 2\gamma, m_{\gamma\gamma} < 110 \text{ MeV}) < 0.6 \times 10^{-8}$  (90% CL).<sup>3</sup> ALAVI-HARATI 99B finds that  $\Gamma(\pi^0 2\gamma, m_{\gamma\gamma} < 240 \text{ MeV}) / \Gamma(\pi^0 2\gamma) = (17.3 \pm 1.3 \pm 1.5)\%$ . Superseded by ABOUZAID 08.<sup>4</sup> BARR 92 find that  $\Gamma(\pi^0 2\gamma, m_{\gamma\gamma} < 240 \text{ MeV}) / \Gamma(\pi^0 2\gamma) < 0.09$  (90% CL).<sup>5</sup> BARR 90c superseded by BARR 92. $\Gamma(\pi^0\gamma e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_{16}/\Gamma$ 

VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.62 ± 0.14 ± 0.09</b>		125	<sup>1</sup> ABOUZAID 07D	KTEV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.34 ± 0.35 ± 0.13	44		ALAVI-HARATI 01E	KTEV	
< 71	90	0	MURAKAMI 99	SPEC	

<sup>1</sup> ABOUZAID 07D includes 1997 (ALAVI-HARATI 01E) and 1999 data. It measures the ratio of  $B(K_L^0 \rightarrow \pi^0\gamma e^+e^-) / B(K_L^0 \rightarrow \pi^0\pi_D^0)$ , where  $\pi_D^0$  is the Dalitz decaying  $\pi^0$ , and uses PDG 06 values  $B(K_L^0 \rightarrow \pi^0\pi^0) = (8.69 \pm 0.04) \times 10^{-4}$ , and  $B(\pi_D^0 \rightarrow e^+e^-\gamma) = (1.198 \pm 0.032) \times 10^{-2}$ . Supersedes ALAVI-HARATI 01E result.Other modes with photons or  $\ell\bar{\ell}$  pairs $\Gamma(2\gamma)/\Gamma_{\text{total}}$  $\Gamma_{17}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.47 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1.
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.54 ± 0.84		<sup>1</sup> BANNER 72B	OSPK	
4.5 ± 1.0	23	ENSTROM 71	OSPK	$K_L^0$ 1.5-9 GeV/c
5.0 ± 1.0		<sup>2</sup> REPELLIN 71	OSPK	
5.5 ± 1.1	90	KUNZ 68	OSPK	Norm. to 3 $\pi(C+N)$

<sup>1</sup> This value uses  $(\eta_{00}/\eta_{+-})^2 = 1.05 \pm 0.14$ . In general,  $\Gamma(2\gamma)/\Gamma_{\text{total}} = [(4.32 \pm 0.55) \times 10^{-4}] [(\eta_{00}/\eta_{+-})^2]$ .<sup>2</sup> Assumes regeneration amplitude in copper at 2 GeV is 22 mb. To evaluate for a given regeneration amplitude and error, multiply by (regeneration amplitude/22mb)<sup>2</sup>.

# Meson Particle Listings

$K_L^0$

$\Gamma(2\gamma)/\Gamma(3\pi^0)$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.802±0.017 OUR FIT</b>				
<b>2.802±0.018 OUR AVERAGE</b>				
2.79 ±0.02 ±0.02	27k	ADINOLFI	03	KLOE
2.81 ±0.01 ±0.02		LAI	03	NA48
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.13 ±0.43	28	BARMIN	71	HLBC
2.24 ±0.28	115	BANNER	69	OSPK
2.5 ±0.7	16	ARNOLD	68B	HLBC Vacuum decay

$\Gamma(2\gamma)/\Gamma(\pi^0\pi^0)$

VALUE	EVTS	DOCUMENT ID	TECN
<b>0.633±0.006 OUR FIT</b>	Error includes scale factor of 1.4.		
<b>0.632±0.004±0.008</b>	110k	BURKHARDT	87 NA31

$\Gamma(3\gamma)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt;7.4 × 10<sup>-8</sup></b>	90	<sup>1</sup> TUNG	11 K391
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<2.4 × 10 <sup>-7</sup>	90	<sup>2</sup> BARR	95c NA31
<sup>1</sup> TUNG 11 reports the result assuming parity violating interaction and using 2005 data (Run-II and III). Assuming parity conserving or phase space interaction, the 90% upper limits obtained are 7.5 × 10 <sup>-8</sup> and 8.6 × 10 <sup>-8</sup> , respectively.			
<sup>2</sup> Assumes a phase-space decay distribution.			

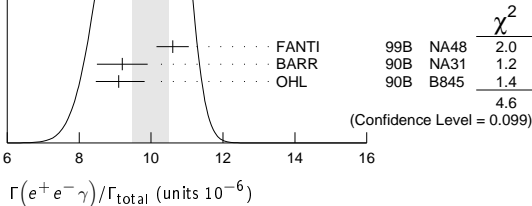
$\Gamma(e^+e^-\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN
<b>9.4±0.4 OUR FIT</b>	Error includes scale factor of 2.0.		
<b>10.0±0.5 OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
10.6±0.2±0.4	6864	<sup>1</sup> FANTI	99B NA48
9.2±0.5±0.5	1053	BARR	90B NA31
9.1±0.4±0.6	919	OHL	90B B845

<sup>1</sup> For FANTI 99B, the ±0.4 systematic error includes for uncertainties in the calculation, primarily uncertainties in the  $\pi^0 \rightarrow e^+e^-\gamma$  and  $K_L^0 \rightarrow \pi^0\pi^0$  branching ratios, evaluated using our 1999 Web edition values.

WEIGHTED AVERAGE  
10.0±0.5 (Error scaled by 1.5)

Values above of weighted average, error, and scale factor are based upon the data in this ideogram only. They are not necessarily the same as our 'best' values, obtained from a least-squares constrained fit utilizing measurements of other (related) quantities as additional information.



$\Gamma(e^+e^-\gamma)/\Gamma(3\pi^0)$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN
<b>4.82±0.21 OUR FIT</b>	Error includes scale factor of 2.0.		
<b>4.63±0.04±0.13</b>	83k	<sup>1</sup> ABOUZAID	07B KTEV

<sup>1</sup> ABOUZAID 07B reports  $[\Gamma(K_L^0 \rightarrow e^+e^-\gamma)/\Gamma(K_L^0 \rightarrow 3\pi^0)] / [3\Gamma(\pi^0 \rightarrow 2\gamma)/\Gamma_{\text{total}} \times \Gamma(\pi^0 \rightarrow e^+e^-\gamma)/\Gamma_{\text{total}}] = (1.3302 \pm 0.0046 \pm 0.0103) \times 10^{-3}$  which we multiply by our best value  $3\Gamma(\pi^0 \rightarrow 2\gamma)/\Gamma_{\text{total}} \times \Gamma(\pi^0 \rightarrow e^+e^-\gamma)/\Gamma_{\text{total}} = 0.0348 \pm 0.0010$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\mu^+\mu^-\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN
<b>3.59±0.11 OUR AVERAGE</b>	Error includes scale factor of 1.3.		
3.62±0.04±0.08	9100	ALAVI-HARATI01G	KTEV
3.4 ±0.6 ±0.4	45	FANTI	97 NA48
3.23±0.23±0.19	197	SPENCER	95 E799

$\Gamma_{17}/\Gamma_6$

$\Gamma_{17}/\Gamma_9$

$\Gamma_{18}/\Gamma$

$\Gamma_{19}/\Gamma$

$\Gamma_{19}/\Gamma_6$

$\Gamma_{20}/\Gamma$

$\Gamma(e^+e^-\gamma\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.95±0.33 OUR AVERAGE</b>				
5.84±0.15±0.32	1543	ALAVI-HARATI01F	KTEV	$E_\gamma^* > 5$ MeV
8.0 ±1.5 <sup>+1.4</sup> <sub>-1.2</sub>	40	SETZU	98 NA31	$E_\gamma^* > 5$ MeV
6.5 ±1.2 ±0.6	58	NAKAYA	94 E799	$E_\gamma^* > 5$ MeV
6.6 ±3.2		MORSE	92 B845	$E_\gamma^* > 5$ MeV

$\Gamma(\mu^+\mu^-\gamma\gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.4<sup>+7.5</sup><sub>-5.9</sub>±0.7</b>	4	ALAVI-HARATI00E	KTEV	$m_{\gamma\gamma} \geq 1$ MeV/c <sup>2</sup>

Charge conjugation × Parity (CP) or Lepton Family number (LF) violating modes, or  $\Delta S = 1$  weak neutral current (S1) modes

$\Gamma(\mu^+\mu^-)/\Gamma(\pi^+\pi^-)$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.48 ±0.05 OUR AVERAGE</b>				
3.474±0.057	6210	AMBROSE	00 B871	
3.87 ±0.30	179	<sup>1</sup> AKAGI	95 SPEC	
3.38 ±0.17	707	HEINSON	95 B791	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.9 ±0.3 ±0.1	178	<sup>2</sup> AKAGI	91B SPEC	In AKAGI 95
3.45 ±0.18 ±0.13	368	<sup>3</sup> HEINSON	91 SPEC	In HEINSON 95
4.1 ±0.5	54	INAGAKI	89 SPEC	In AKAGI 91B
2.8 ±0.3 ±0.2	87	MATHIAZHA...	89B SPEC	In HEINSON 91

<sup>1</sup> AKAGI 95 gives this number multiplied by the PDG 1992 average for  $\Gamma(K_L^0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$ .

<sup>2</sup> AKAGI 91B give this number multiplied by the 1990 PDG average for  $\Gamma(K_L^0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$ .

<sup>3</sup> HEINSON 91 give  $\Gamma(K_L^0 \rightarrow \mu\mu)/\Gamma_{\text{total}}$ . We divide out the  $\Gamma(K_L^0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$  PDG average which they used.

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-10}$ )	CL%	EVTS	DOCUMENT ID	TECN
<b>0.087±0.057</b>				
<b>-0.041</b>		4	AMBROSE	98 B871

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.6	90	1	AKAGI	95 SPEC
<0.41	90	0	<sup>1</sup> ARISAKA	93B B791

<sup>1</sup> ARISAKA 93B includes all events with <6 MeV radiated energy.

$\Gamma(\pi^+\pi^-\pi^+e^+e^-)/\Gamma_{\text{total}}$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.11±0.19 OUR AVERAGE</b>					
3.08±0.09±0.18		1125	<sup>1</sup> LAI	03c NA48	
3.2 ±0.6 ±0.4		37	ADAMS	98 KTEV	
4.4 ±1.3 ±0.5		13	TAKEUCHI	98 SPEC	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<4.6 90 NOMURA 97 SPEC  $m_{ee} > 4$  MeV

<sup>1</sup> LAI 03c second error is 0.15(syst)±0.10(norm) combined in quadrature. The normalization uses  $\text{BR}(K_L \rightarrow \pi^+\pi^-\pi^0) * \text{BR}(\pi^0 \rightarrow e^+e^-) = (1.505 \pm 0.047) \times 10^{-3}$  from our 2000 Edition.

$\Gamma(\pi^0\pi^0\pi^+e^+e^-)/\Gamma_{\text{total}}$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN
<b>&lt;6.6</b>		90	1	ALAVI-HARATI02C E799

$\Gamma(\pi^0\pi^0\mu^+\mu^-)/\Gamma_{\text{total}}$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt;9.2 × 10<sup>-11</sup></b>	90	<sup>1</sup> ABOUZAID	11A E799

<sup>1</sup> ABOUZAID 11A also reports  $\text{B}(K_L^0 \rightarrow \pi^0\pi^0X^0 \rightarrow \pi^0\pi^0\mu^+\mu^-) < 1.0 \times 10^{-10}$  at 90% C.L., where the  $X^0$  is a possible new neutral boson that was reported by PARK 05 with a mass of  $214.3 \pm 0.5$  MeV/c<sup>2</sup>.

$\Gamma(\mu^+\mu^-\pi^+e^+e^-)/\Gamma_{\text{total}}$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.69±0.27 OUR AVERAGE</b>					
2.69±0.24±0.12		131	<sup>1</sup> ALAVI-HARATI03B	KTEV	
2.9 <sup>+6.7</sup> <sub>-2.4</sub>		1	GU	96 E799	

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.62±0.40±0.17 43 ALAVI-HARATI01H KTEV Sup. by ALAVI-HARATI 03B

<4900 90 BALATS 83 SPEC

<sup>1</sup> ALAVI-HARATI 03B also measures the linear slope  $\alpha = -1.59 \pm 0.37$ .

$\Gamma(e^+e^-e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{29}/\Gamma$ Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.56 \pm 0.21</math></b>					<b>OUR AVERAGE</b>
$3.30 \pm 0.24 \pm 0.25$		200	<sup>1</sup> LAI	05B	NA48
$3.72 \pm 0.18 \pm 0.23$		441	ALAVI-HARATI	01D	KTEV
$3.96 \pm 0.78 \pm 0.32$		27	GU	94	E799
$3.07 \pm 1.25 \pm 0.26$		6	VAGINS	93	B845
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$6 \pm 2 \pm 1$		18	<sup>2</sup> AKAGI	95	SPEC $m_{ee} > 470$ MeV
$7 \pm 3 \pm 2$		6	<sup>2</sup> AKAGI	95	SPEC $m_{ee} > 470$ MeV
$10.4 \pm 3.7 \pm 1.1$		8	<sup>3</sup> BARR	95	NA31
$6 \pm 2 \pm 1$		18	AKAGI	93	CNTR Sup. by AKAGI 95
$4 \pm 3$		2	BARR	91	NA31 Sup. by BARR 95

<sup>1</sup> LAI 05B uses 1998 and 1999 data. Data are normalized to the observed events of  $K_L^0 \rightarrow \pi^+\pi^-\pi^0$  ( $\pi^0$  into Dalitz pair) and PDG 04 values are used for  $B(K_L^0 \rightarrow \pi^+\pi^-\pi^0)$  and  $B(\pi^0 \rightarrow e^+e^-\gamma)$ . The systematic error includes a normalization error of  $\pm 0.10$ .<sup>2</sup> Values are for the total branching fraction, acceptance-corrected for the  $m_{ee}$  cuts shown.<sup>3</sup> Distribution of angles between two  $e^+e^-$  pair planes favors  $CP = -1$  for  $K_L^0$ . $\Gamma(\pi^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{30}/\Gamma$ Violates  $CP$  in leading order. Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 0.38</math></b>		90	ALAVI-HARATI	00D	KTEV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 5.1$		90	0	HARRIS	93 E799

 $\Gamma(\pi^0e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{31}/\Gamma$ Violates  $CP$  in leading order. Direct and indirect  $CP$ -violating contributions are expected to be comparable and to dominate the  $CP$ -conserving part. LAI 02B result suggests that  $CP$ -violation effects dominate. Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE (units $10^{-10}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.8</math></b>		90	<sup>1</sup> ALAVI-HARATI	04A	KTEV combined result
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 3.5$		90	ALAVI-HARATI	04A	KTEV
$0.0047 \pm 0.0022$ $-0.0018$			<sup>2</sup> LAI	02B	NA48 $CP$ -conserving part
$< 5.1$		90	2	ALAVI-HARATI 01	KTEV
$0.01$ to $0.02$			ALAVI-HARATI	99B	KTEV $CP$ -conserving part
$< 43$		90	0	HARRIS	93B E799
$< 75$		90	0	BARKER	90 E731
$< 55$		90	0	OHL	90 B845
$< 400$		90		BARR	88 NA31
$< 3200$		90		JASTRZEM...	88 SPEC

<sup>1</sup> Combined result of ALAVI-HARATI 04A 1999-2000 data set and ALAVI-HARATI 01 1997 data set.<sup>2</sup> LAI 02B uses the absence of a signal in  $K_L^0 \rightarrow \pi^0\gamma\gamma$  with  $m(\gamma\gamma) < m(\pi^0)$  and their  $a_V$  value to predict this value. $\Gamma(\pi^0\nu\bar{\nu})/\Gamma_{\text{total}}$   $\Gamma_{32}/\Gamma$ Violates  $CP$  in leading order. Test of direct  $CP$  violation since the indirect  $CP$ -violating and  $CP$ -conserving contributions are expected to be suppressed. Test of  $\Delta S = 1$  weak neutral current.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 0.26</math></b>		90	<sup>1</sup> AHN	10 K391
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.67$		90	<sup>2</sup> AHN	08 K391
$< 2.1$		90	<sup>3</sup> AHN	06 K391
$< 5.9$		90	ALAVI-HARATI	00 KTEV
$< 16$		90	ADAMS	99 KTEV
$< 580$		90	WEAVER	94 E799
$< 2200$		90	GRAHAM	92 CNTR

<sup>1</sup> Obtained combining Run-2 (AHN 08) and Run-3 data.<sup>2</sup> Value obtained using data from February to April 2005.<sup>3</sup> Value obtained analyzing 10% of data of RUN 1 (performed in 2004). $\Gamma(\pi^0\pi^0\nu\bar{\nu})/\Gamma_{\text{total}}$   $\Gamma_{33}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 8.1 \times 10^{-7}</math></b>		90	<sup>1</sup> OGATA	11 K391
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.7 \times 10^{-5}$		90	<sup>2</sup> NIX	07 K391

<sup>1</sup> Using 2005 Run-I data. OGATA 11 also sets a limit on the  $K_L^0 \rightarrow \pi^0\pi^0 X \rightarrow$  invisible particles process: the limit on the branching fraction varied from  $7.0 \times 10^{-7}$  to  $4.0 \times 10^{-5}$  for the mass of  $X$  ranging from 50 to 200 MeV/ $c^2$ .<sup>2</sup> Observed 1 event with expected background of  $0.43 \pm 0.35$  events. NIX 07 also measured  $B(K_L^0 \rightarrow \pi^0\pi^0 P) < 1.2 \times 10^{-6}$  at 90% CL, where  $P$  is the pseudoscalar particle and  $m_P < 100$  MeV. $\Gamma(e^\pm\mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{34}/\Gamma$ 

Test of lepton family number conservation.

VALUE (units $10^{-11}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 0.47</math></b>		90	AMBROSE	98B	B871
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 9.4$		90	0	AKAGI	95 SPEC
$< 3.9$		90	0	ARISA KA	93 B791
$< 3.3$		90	0	<sup>1</sup> ARISA KA	93 B791

<sup>1</sup> This is the combined result of ARISA KA 93 and MATHIAZHAGAN 89. $\Gamma(e^\pm e^\pm\mu^\mp\mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{35}/\Gamma$ 

Test of lepton family number conservation.

VALUE (units $10^{-11}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.12</math></b>		90	0	ALAVI-HARATI	03B KTEV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 12.3$		90	0	<sup>1</sup> ALAVI-HARATI	01H KTEV Sup. by ALAVI-HARATI 03B
$< 610$		90	0	<sup>1</sup> GU	96 E799

<sup>1</sup> Assuming uniform phase space distribution. $\Gamma(\pi^0\mu^\pm e^\mp)/\Gamma_{\text{total}}$   $\Gamma_{36}/\Gamma$ 

Test of lepton family number conservation.

VALUE (units $10^{-10}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 0.76</math></b>		90	ABOUZAID	08c KTEV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 62$		90	ARISA KA	98 E799

 $\Gamma(\pi^0\pi^0\mu^\pm e^\mp)/\Gamma_{\text{total}}$   $\Gamma_{37}/\Gamma$ 

Test of lepton family number conservation.

VALUE (units $10^{-10}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.7</math></b>		90	ABOUZAID	08c KTEV

 $V_{ud}, V_{us}$ , THE CABIBBO ANGLE, AND CKM UNITARITY

Updated May 2016 by E. Blucher (Univ. of Chicago) and W.J. Marciano (BNL)

The Cabibbo-Kobayashi-Maskawa (CKM) [1,2] three-generation quark mixing matrix written in terms of the Wolfenstein parameters  $(\lambda, A, \rho, \eta)$  [3] nicely illustrates the orthonormality constraint of unitarity and central role played by  $\lambda$ .

$$V_{\text{CKM}} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} = \begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix} + \mathcal{O}(\lambda^4). \quad (1)$$

That cornerstone is a carryover from the two-generation Cabibbo angle,  $\lambda = \sin(\theta_{\text{Cabibbo}}) = V_{us}$ . Its value is a critical ingredient in determinations of the other parameters and in tests of CKM unitarity.

Until about 11 years ago, the precise value of  $\lambda$  was somewhat controversial, with kaon decays suggesting [4]  $\lambda \simeq 0.220$ , while indirect determinations via nuclear  $\beta$ -decays implied a somewhat larger  $\lambda \simeq 0.225 - 0.230$ . This difference resulted in a 2 – 2.5 sigma deviation from the unitarity requirement

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1, \quad (2)$$

a potential signal [5] for new physics effects. Below, we discuss the current status of  $V_{ud}$ ,  $V_{us}$ , and their associated unitarity test in Eq. (2). (Since  $|V_{ub}|^2 \simeq 1 \times 10^{-5}$  is negligibly small,

# Meson Particle Listings

$K_L^0$

it is ignored in this discussion.) Eq. (2) is currently the most stringent test of unitarity in the CKM matrix.

## $V_{ud}$

The value of  $V_{ud}$  has been obtained from superallowed nuclear, neutron, and pion decays. Currently, the most precise determination of  $V_{ud}$  comes from a set of superallowed nuclear beta-decays [5] ( $0^+ \rightarrow 0^+$  transitions). Measuring their half-lives,  $t$ , and  $Q$  values that give the decay rate factor,  $f$ , leads to a precise determination of  $V_{ud}$  via the master formula [6–10]

$$|V_{ud}|^2 = \frac{2984.48(5) \text{ sec}}{ft(1 + \Delta)}, \quad (3)$$

where  $\Delta$  denotes the entire effect of electroweak radiative corrections (RC), nuclear structure, and isospin violating nuclear effects.  $\Delta$  is nucleus-dependent, ranging from about +3.0% to +3.6% for the best measured superallowed decays.

The most recent analysis of 14 precisely measured superallowed transitions by Hardy and Towner [11] gives a weighted average of

$$V_{ud} = 0.97417(5)_{\text{exp.}}(9)_{\text{nucl.dep.}}(18)_{\text{RC}} \text{ (superallowed)}, \quad (4)$$

which, assuming unitarity, corresponds to  $\lambda = 0.2258(9)$ . This recent determination of  $V_{ud}$  has shifted downward compared to the 2014 value of 0.97425(22) primarily from improvements in the nuclear isospin breaking corrections [11]. It is now closer to the central value quoted in 2007.

Combined measurements of the neutron lifetime,  $\tau_n$ , and the ratio of axial-vector/vector couplings,  $g_A \equiv G_A/G_V$ , via neutron decay asymmetries can also be used to determine  $V_{ud}$ :

$$|V_{ud}|^2 = \frac{4908.7(1.9) \text{ sec}}{\tau_n(1 + 3g_A^2)}, \quad (5)$$

where the error stems from uncertainties in the electroweak radiative corrections [7] due to hadronic loop effects. Those effects were updated and their error was reduced by about a factor of 2 [8], leading to a  $\pm 0.0002$  theoretical uncertainty in  $V_{ud}$  (common to all  $V_{ud}$  extractions). Using the world averages from this *Review*

$$\begin{aligned} \tau_n^{\text{ave}} &= 880.3(1.1) \text{ sec} \quad (\times 1.9 \text{ PDG scale factor}) \\ g_A^{\text{ave}} &= 1.2723(23) \quad (\times 2.2 \text{ PDG scale factor}) \end{aligned} \quad (6)$$

leads to

$$V_{ud} = 0.9758(6)_{\tau_n}(15)_{g_A}(2)_{\text{RC}}, \quad (7)$$

with the error dominated by  $g_A$  uncertainties. We note that the larger  $g_A$  now adopted in Eq. (6) leads to a value of  $V_{ud}$  that is still somewhat high, but in accord with the superallowed nuclear beta decay result in Eq. (4). Future neutron studies [12] are expected to resolve any current inconsistencies and significantly

reduce the uncertainties in  $g_A$  and  $\tau_n$ , potentially making them a competitive way to determine  $V_{ud}$  without nuclear physics uncertainties.

The PIBETA experiment at PSI measured the very small ( $\mathcal{O}(10^{-8})$ ) branching ratio for  $\pi^+ \rightarrow \pi^0 e^+ \nu_e$  with about  $\pm 1/2\%$  precision. Their result gives [13]

$$V_{ud} = 0.9749(26) \left[ \frac{BR(\pi^+ \rightarrow e^+ \nu_e(\gamma))}{1.2352 \times 10^{-4}} \right]^{\frac{1}{2}} \quad (8)$$

which is normalized using the very precisely determined theoretical prediction for  $BR(\pi^+ \rightarrow e^+ \nu_e(\gamma)) = 1.2352(5) \times 10^{-4}$  [6], rather than the experimental branching ratio from this *Review* of  $1.230(4) \times 10^{-4}$  which would lower the value to  $V_{ud} = 0.9728(30)$ . Theoretical uncertainties in the pion  $\beta$ -decay determination are very small; however, much higher statistics would be required to make this approach competitive with others.

## $V_{us}$

$|V_{us}|$  may be determined from kaon decays, hyperon decays, and tau decays. Previous determinations have most often used  $K\ell 3$  decays:

$$\Gamma_{K\ell 3} = \frac{G_F^2 M_K^5}{192\pi^3} S_{EW}(1 + \delta_K^\ell + \delta_{SU2}) C^2 |V_{us}|^2 f_+^2(0) I_K^\ell. \quad (9)$$

Here,  $\ell$  refers to either  $e$  or  $\mu$ ,  $G_F$  is the Fermi constant,  $M_K$  is the kaon mass,  $S_{EW}$  is the short-distance radiative correction,  $\delta_K^\ell$  is the mode-dependent long-distance radiative correction,  $f_+(0)$  is the calculated form factor at zero momentum transfer for the  $\ell\nu$  system, and  $I_K^\ell$  is the phase-space integral, which depends on measured semileptonic form factors. For charged kaon decays,  $\delta_{SU2}$  is the deviation from one of the ratio of  $f_+(0)$  for the charged to neutral kaon decay; it is zero for the neutral kaon.  $C^2$  is 1 (1/2) for neutral (charged) kaon decays. Most early determinations of  $|V_{us}|$  were based solely on  $K \rightarrow \pi e \nu$  decays;  $K \rightarrow \pi \mu \nu$  decays were not used because of large uncertainties in  $I_K^\mu$ . The experimental measurements are the semileptonic decay widths (based on the semileptonic branching fractions and lifetime) and form factors (allowing calculation of the phase space integrals). Theory is needed for  $S_{EW}$ ,  $\delta_K^\ell$ ,  $\delta_{SU2}$ , and  $f_+(0)$ .

Many measurements during the last decade have resulted in a significant shift in  $V_{us}$ . Most importantly, recent measurements of the  $K \rightarrow \pi e \nu$  branching fractions are significantly different than earlier PDG averages, probably as a result of inadequate treatment of radiation in older experiments. This effect was first observed by BNL E865 [14] in the charged kaon system and then by KTeV [15,16] in the neutral kaon system; subsequent measurements were made by KLOE [17–20], NA48 [21–23], and ISTRA+ [24]. Current averages (*e.g.*, by the PDG [25] or Flavianet [26]) of the semileptonic branching fractions are based only on recent, high-statistics experiments

where the treatment of radiation is clear. In addition to measurements of branching fractions, new measurements of lifetimes [27] and form factors [28–32], have resulted in improved precision for all of the experimental inputs to  $V_{us}$ . Precise measurements of form factors for  $K_{\mu 3}$  decay make it possible to use both semileptonic decay modes to extract  $V_{us}$ .

Following the analysis of Moulson [33] and the Flavianet group [26], one finds, after including the isospin violating up-down mass difference effect, the values of  $|V_{us}|f_+(0)$  in Table 1. The average of these measurements gives

$$f_+(0)|V_{us}| = 0.2165(4). \quad (10)$$

Figure 1 shows a comparison of these results with the PDG evaluation from 2002 [34], as well as  $f_+(0)(1-|V_{ud}|^2-|V_{ub}|^2)^{1/2}$ , the expectation for  $f_+(0)|V_{us}|$  assuming unitarity, based on  $|V_{ud}| = 0.97417 \pm 0.00021$ , and  $|V_{ub}| = (4.1 \pm 0.4) \times 10^{-3}$  [35].

Lattice calculations of  $f_+(0)$  have been carried out for 2, 2+1, and 2+1+1 quark flavors and range from about 0.96 to 0.97. Here, we use  $f_+(0) = 0.9677(37)$ , the 2015 preliminary (2+1)-flavor FLAG average reported by Rosner, Stone, and Van de Water in footnote 10 of their PDG review of pseudoscalar decay constants [35], in Eq. (10), and find

$$|V_{us}| = \lambda = 0.2237(4)_{\text{exp+RC}}(9)_{\text{lattice}} (K_{\ell 3} \text{ Decays}). \quad (11)$$

**Table 1:**  $|V_{us}|f_+(0)$  from  $K_{\ell 3}$ .

Decay Mode	$ V_{us} f_+(0)$
$K^\pm e 3$	$0.2172 \pm 0.0008$
$K^\pm \mu 3$	$0.2170 \pm 0.0011$
$K_L e 3$	$0.2163 \pm 0.0006$
$K_L \mu 3$	$0.2166 \pm 0.0006$
$K_S e 3$	$0.2155 \pm 0.0013$
Average	$0.2165 \pm 0.0004$

A value of  $V_{us}$  can also be obtained from a comparison of the radiative inclusive decay rates for  $K \rightarrow \mu\nu(\gamma)$  and  $\pi \rightarrow \mu\nu(\gamma)$  combined with a lattice gauge theory calculation of  $f_{K^+}/f_{\pi^+}$  via [42]

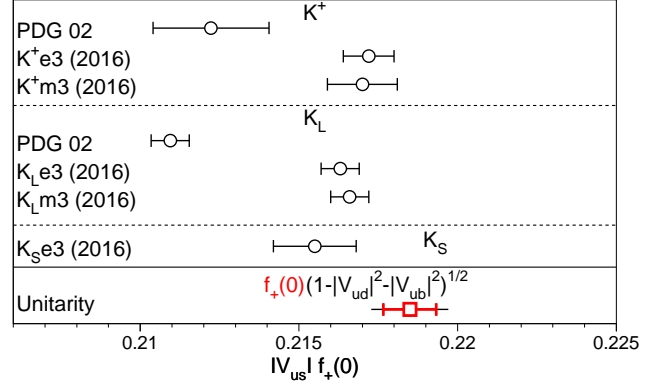
$$\frac{|V_{us}|f_{K^+}}{|V_{ud}|f_{\pi^+}} = 0.23871(20) \left[ \frac{\Gamma(K \rightarrow \mu\nu(\gamma))}{\Gamma(\pi \rightarrow \mu\nu(\gamma))} \right]^{\frac{1}{2}} \quad (12)$$

with the small error coming from electroweak radiative corrections and isospin breaking effects. Employing

$$\frac{\Gamma(K \rightarrow \mu\nu(\gamma))}{\Gamma(\pi \rightarrow \mu\nu(\gamma))} = 1.3367(29), \quad (13)$$

which includes the recent update  $\Gamma(K \rightarrow \mu\nu(\gamma)) = 5.134(11) \times 10^7 \text{s}^{-1}$  [33,43] and [35]

$$f_{K^+}/f_{\pi^+} = 1.1928(26) \quad (14)$$



**Figure 1:** Comparison of determinations of  $|V_{us}|f_+(0)$  from this review (labeled 2016), from the PDG 2002, and with the prediction from unitarity using  $|V_{ud}|$  and the lattice calculation of  $f_+(0) = 0.9677(37)$  [35]. For  $f_+(0)(1-|V_{ud}|^2-|V_{ub}|^2)^{1/2}$ , the inner error bars are from the quoted uncertainty in  $f_+(0)$ ; the total uncertainties include the  $|V_{ud}|$  and  $|V_{ub}|$  errors.

along with the value of  $V_{ud}$  in Eq. (4) leads to

$$|V_{us}| = 0.22540(53)_{\text{exp}}(19)_{\text{RC}}(49)_{\text{lattice}} (K_{\mu 2} \text{ Decays}). \quad (15)$$

Together, a weighted average of the  $K_{\ell 3}$  (Eq. (11)) and  $K_{\mu 2}$  (Eq. (15)) results gives

$$|V_{us}| = 0.2248(6). \quad (16)$$

It should be mentioned that hyperon decay fits suggest [45]

$$|V_{us}| = 0.2250(27) \text{ (Hyperon Decays)} \quad (17)$$

modulo SU(3) breaking effects that could shift that value up or down. We note that a representative effort [46] that incorporates SU(3) breaking found  $V_{us} = 0.226(5)$ . Strangeness changing tau decays, averaging both inclusive and exclusive measurements, currently give [47]

$$|V_{us}| = 0.2202(15) \text{ (Tau Decays)}, \quad (18)$$

which differs by about 3 sigma from the kaon determination discussed above, and would, if combined with  $V_{ud}$  from super-allowed beta decays, lead to a 2.6 sigma deviation from unitarity. This discrepancy results mainly from the inclusive tau decay results that rely on Finite Energy Sum Rule techniques and assumptions. Further investigation of that approach seems to be warranted.

Employing the values of  $V_{ud}$  and  $V_{us}$  from Eq. (4) and Eq. (16), respectively, leads to the unitarity consistency check

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9995(4)(3). \quad (19)$$

where the first error is the uncertainty from  $|V_{ud}|^2$  and the second error is the uncertainty from  $|V_{us}|^2$ .

## CKM Unitarity Constraints

The current good experimental agreement with unitarity,  $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9995(5)$ , provides strong confirmation of Standard Model radiative corrections (which range between 3-4% depending on the nucleus used) at better than the 50 sigma level [48]. In addition, it implies constraints on “New Physics” effects at both the tree and quantum loop levels. Those effects could be in the form of contributions to nuclear beta decays,  $K$  decays and/or muon decays, with the last of these providing normalization via the muon lifetime [49], which is used to obtain the Fermi constant,  $G_\mu = 1.1663787(6) \times 10^{-5} \text{GeV}^{-2}$ .

In the following sections, we illustrate the implications of CKM unitarity for (1) exotic muon decays [50] (beyond ordinary muon decay  $\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$ ) and (2) new heavy quark mixing  $V_{uD}$  [51]. Other examples in the literature [52,53] include  $Z_\chi$  boson quantum loop effects, supersymmetry, leptoquarks, compositeness etc.

## Exotic Muon Decays

If additional lepton flavor violating decays such as  $\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_\mu$  (wrong neutrinos) occur, they would cause confusion in searches for neutrino oscillations at, for example, muon storage rings/neutrino factories or other neutrino sources from muon decays. Calling the rate for all such decays  $\Gamma(\text{exotic } \mu \text{ decays})$ , they should be subtracted before the extraction of  $G_\mu$  and normalization of the CKM matrix. Since that is not done and unitarity works, one has (at one-sided 95% CL)

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 - BR(\text{exotic } \mu \text{ decays}) \geq 0.9987 \quad (20)$$

or

$$BR(\text{exotic } \mu \text{ decays}) \leq 0.0013. \quad (21)$$

This bound is a factor of 10 better than the direct experimental bound on  $\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_\mu$ .

## New Heavy Quark Mixing

Heavy  $D$  quarks naturally occur in fourth quark generation models and some heavy quark “new physics” scenarios such as  $E_6$  grand unification. Their mixing with ordinary quarks gives rise to  $V_{uD}$  which is constrained by unitarity (one sided 95% CL)

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 - |V_{uD}|^2 \geq 0.9987 \\ |V_{uD}| \leq 0.04. \quad (22)$$

A similar constraint applies to heavy neutrino mixing and the couplings  $V_{\mu N}$  and  $V_{eN}$ .

## References

1. N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
2. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
3. L. Wolfenstein, Phys. Rev. Lett. **51**, 1945 (1983).
4. S. Eidelman *et al.* [Particle Data Group], Phys. Lett. **B592**, 1 (2004).
5. J.C. Hardy and I.S. Towner, Rep. Prog. Phys. **73**, 046301 (2010).
6. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **71**, 3629 (1993).
7. A. Czarnecki, W.J. Marciano, and A. Sirlin, Phys. Rev. **D70**, 093006 (2004) [hep-ph/0406324].
8. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **96**, 032002 (2006) [hep-ph/0510099].
9. J.C. Hardy and I.S. Towner, Phys. Rev. **C77**, 025501 (2008).
10. J.C. Hardy and I.S. Towner, Phys. Rev. **C79**, 055502 (2009).
11. J.C. Hardy and I.S. Towner, Phys. Rev. **C91**, 0255012 (2015).
12. H. Abele, Prog. in Part. Nucl. Phys. **60**, 1 (2008).
13. D. Pocanic *et al.*, Phys. Rev. Lett. **93**, 181803 (2004) [hep-ex/0312030].
14. A. Sher *et al.*, Phys. Rev. Lett. **91**, 261802 (2003).
15. T. Alexopoulos *et al.* [KTeV Collab.], Phys. Rev. Lett. **93**, 181802 (2004) [hep-ex/0406001].
16. T. Alexopoulos *et al.* [KTeV Collab.], Phys. Rev. **D70**, 092006 (2004) [hep-ex/0406002].
17. F. Ambrosino *et al.* [KLOE Collab.], Phys. Lett. **B632**, 43 (2006) [hep-ex/0508027].
18. F. Ambrosino *et al.* [KLOE Collab.], Phys. Lett. **B638**, 140 (2006) [hep-ex/0603041].
19. F. Ambrosino *et al.* [KLOE Collab.], Phys. Lett. **B636**, 173 (2006) [hep-ex/0601026].
20. F. Ambrosino *et al.* [KLOE Collab.], PoS **HEP2005**, 287 (2006) [hep-ex/0510028].
21. A. Lai *et al.* [NA48 Collab.], Phys. Lett. **B602**, 41 (2004) [hep-ex/0410059].
22. A. Lai *et al.* [NA48 Collab.], Phys. Lett. **B645**, 26 (2007) [hep-ex/0611052].
23. J.R. Batley *et al.* [NA48/2 Collab.], Eur. Phys. J. **C50**, 329 (2007) [hep-ex/0702015].
24. V.I. Romanovsky *et al.*, [hep-ex/0704.2052].
25. K.A. Olive *et al.* [Particle Data Group], Chin. Phys. C **38**, 090001 (2014).
26. Flavianet Working Group on Precise SM Tests in K Decays, <http://www.lnf.infn.it/wg/vus>; M. Antonelli *et al.*, Eur. Phys. J. **C69**, 399 (2010). For a recent detailed review, see M. Antonelli *et al.*, [hep-ph/0907.5386].
27. F. Ambrosino *et al.* [KLOE Collab.], Phys. Lett. **B626**, 15 (2005) [hep-ex/0507088].
28. T. Alexopoulos *et al.* [KTeV Collab.], Phys. Rev. **D70**, 092007 (2004) [hep-ex/0406003].
29. E. Abouzaid *et al.* [KTeV Collab.], Phys. Rev. **D74**, 097101 (2006) [hep-ex/0608058].
30. F. Ambrosino *et al.* [KLOE Collab.], Phys. Lett. **B636**, 166 (2006) [hep-ex/0601038].
31. A. Lai *et al.* [NA48 Collab.], Phys. Lett. **B604**, 1 (2004) [hep-ex/0410065].
32. O.P. Yushchenko *et al.*, Phys. Lett. **B589**, 111 (2004) [hep-ex/0404030].
33. M. Moulson, [hep-ex/1411.5251].

34. K. Hagiwara *et al.* [Particle Data Group], Phys. Rev. **D66**, 1 (2002).
35. J. Rosner, S. Stone, R. Van de Water, “*Leptonic decays of charged pseudoscalar mesons*”, in this *Review*.
36. A. Bazavov *et al.*, Phys. Rev. Lett. **112**, 112001 (2014).
37. H. Leutwyler and M. Roos, Z. Phys. **C25**, 91 (1984).
38. D. Becirevic *et al.*, Nucl. Phys. **B705**, 339 (2005) [hep-ph/0403217].
39. J. Bijnens and P. Talavera, Nucl. Phys. **B669**, 341 (2003).
40. V. Cirigliano *et al.*, JHEP **0504**, 006 (2005) [hep-ph/0503108].
41. M. Jamin, J.A. Oller, and A. Pich, JHEP **02**, 047 (2004).
42. V. Cirigliano and H. Neufeld, Phys. Lett. **B700**, 7 (2011); W.J. Marciano, Phys. Rev. Lett. **93**, 231803 (2004) [hep-ph/0402299].
43. D. Babusci *et al.* [KLOE Collab.], Phys. Lett. **B738**, 128 (2014) [arXiv:1407.2028].
44. R.J. Dowdall *et al.*, Phys. Rev. D **88**, 074504 (2013).
45. N. Cabibbo, E.C. Swallow, and R. Winston, Phys. Rev. Lett. **92**, 251803 (2004) [hep-ph/0307214].
46. V. Mateu and A. Pich, JHEP **0510**, 041 (2005) [hep-ph/0509045].
47. I. Nugent, [arXiv:1301.0637]; Y. Amhis *et al.* [Heavy Flavor Averaging Group], [arXiv:1207.1158].
48. A. Sirlin, Rev. Mod. Phys. **50**, 573 (1978).
49. D. Webber *et al.* [MuLan Collab.], Phys. Rev. Lett. **106**, 041803 (2011); V. Tishchenko *et al.* [MuLan Collab.], Phys. Rev. **D87**, 052003 (2013).
50. K.S. Babu and S. Pakvasa, hep-ph/0204236.
51. W. Marciano and A. Sirlin, Phys. Rev. Lett. **56**, 22 (1986); P. Langacker and D. London, Phys. Rev. **D38**, 886 (1988).
52. W. Marciano and A. Sirlin, Phys. Rev. **D35**, 1672 (1987).
53. R. Barbieri *et al.*, Phys. Lett. **156B**, 348 (1985); K. Hagiwara *et al.*, Phys. Rev. Lett. **75**, 3605 (1995); A. Kurylov and M. Ramsey-Musolf, Phys. Rev. Lett. **88**, 071804 (2000).

ENERGY DEPENDENCE OF  $K_L^0$  DALITZ PLOT

For discussion, see note on Dalitz plot parameters in the  $K^\pm$  section of the Particle Listings above. For definitions of  $a_V$ ,  $a_T$ ,  $a_H$ , and  $a_Y$ , see the earlier version of the same note in the 1982 edition of this *Review* published in Physics Letters **111B** 70 (1982).

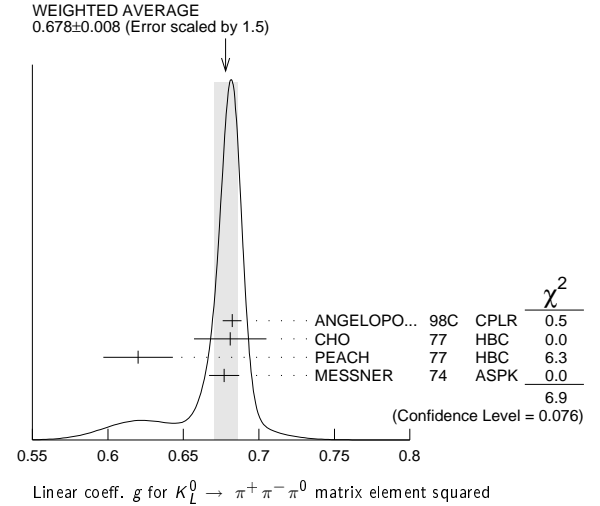
$$|\text{matrix element}|^2 = 1 + gu + hu^2 + jv + kv^2 + fuv$$

where  $u = (s_3 - s_0) / m_\pi^2$  and  $v = (s_2 - s_1) / m_\pi^2$

LINEAR COEFFICIENT  $g$  FOR  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.678 \pm 0.008</math> OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
$0.6823 \pm 0.0044 \pm 0.0044$	500k	ANGELOPO...	98c CPLR	
$0.681 \pm 0.024$	6499	CHO	77 HBC	
$0.620 \pm 0.023$	4709	PEACH	77 HBC	
$0.677 \pm 0.010$	509k	MESSNER	74 ASPK	$a_Y = -0.917 \pm 0.013$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.69 \pm 0.07$	192	<sup>1</sup> BALDO...	75 HLBC	
$0.590 \pm 0.022$	56k	<sup>1</sup> BUCHANAN	75 SPEC	$a_H = -0.277 \pm 0.010$
$0.619 \pm 0.027$	20k	<sup>1,2</sup> BISI	74 ASPK	$a_T = -0.282 \pm 0.011$
$0.612 \pm 0.032$		<sup>1</sup> ALEXANDER	73B HBC	
$0.73 \pm 0.04$	3200	<sup>1</sup> BRANDENB...	73 HBC	
$0.608 \pm 0.043$	1486	<sup>1</sup> KRENZ	72 HLBC	$a_T = -0.277 \pm 0.018$
$0.650 \pm 0.012$	29k	<sup>1</sup> ALBROW	70 ASPK	$a_Y = -0.858 \pm 0.015$
$0.593 \pm 0.022$	36k	<sup>1,3</sup> BUCHANAN	70 SPEC	$a_H = -0.278 \pm 0.010$
$0.664 \pm 0.056$	4400	<sup>1</sup> SMITH	70 OSPK	$a_T = -0.306 \pm 0.024$
$0.400 \pm 0.045$	2446	<sup>1</sup> BASILE	68B OSPK	$a_T = -0.188 \pm 0.020$
$0.649 \pm 0.044$	1350	<sup>1</sup> HOPKINS	67 HBC	$a_T = -0.294 \pm 0.018$
$0.428 \pm 0.055$	1198	<sup>1</sup> NEFKENS	67 OSPK	$a_H = -0.204 \pm 0.025$

- <sup>1</sup> Quadratic dependence required by some experiments. (See sections on “QUADRATIC COEFFICIENT  $h$ ” and “QUADRATIC COEFFICIENT  $k$ ” below.) Correlations prevent us from averaging results of fits not including  $g$ ,  $h$ , and  $k$  terms.
- <sup>2</sup> BISI 74 value comes from quadratic fit with quad. term consistent with zero.  $g$  error is thus larger than if linear fit were used.
- <sup>3</sup> BUCHANAN 70 result revised by BUCHANAN 75 to include radiative correlations and to use more reliable  $K_L^0$  momentum spectrum of second experiment (had same beam).

QUADRATIC COEFFICIENT  $h$  FOR  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$ 

VALUE	EVTS	DOCUMENT ID	TECN
<b><math>0.076 \pm 0.006</math> OUR AVERAGE</b>			
$0.061 \pm 0.004 \pm 0.015$	500k	ANGELOPO...	98C CPLR
$0.095 \pm 0.032$	6499	CHO	77 HBC
$0.048 \pm 0.036$	4709	PEACH	77 HBC
$0.079 \pm 0.007$	509k	MESSNER	74 ASPK
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.011 \pm 0.018$	29k	<sup>1</sup> ALBROW	70 ASPK
$0.043 \pm 0.052$	4400	<sup>1</sup> SMITH	70 OSPK

See notes in section “LINEAR COEFFICIENT  $g$  FOR  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$  | MATRIX ELEMENT|<sup>2</sup>” above.

- <sup>1</sup> Quadratic coefficients  $h$  and  $k$  required by some experiments. (See section on “QUADRATIC COEFFICIENT  $k$ ” below.) Correlations prevent us from averaging results of fits not including  $g$ ,  $h$ , and  $k$  terms.

QUADRATIC COEFFICIENT  $k$  FOR  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$ 

VALUE	EVTS	DOCUMENT ID	TECN
<b><math>0.0099 \pm 0.0015</math> OUR AVERAGE</b>			
$0.0104 \pm 0.0017 \pm 0.0024$	500k	ANGELOPO...	98C CPLR
$0.024 \pm 0.010$	6499	CHO	77 HBC
$-0.008 \pm 0.012$	4709	PEACH	77 HBC
$0.0097 \pm 0.0018$	509k	MESSNER	74 ASPK

LINEAR COEFFICIENT  $j$  FOR  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$  (CP-VIOLATING TERM)

Listed in CP-violation section below.

QUADRATIC COEFFICIENT  $f$  FOR  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$  (CP-VIOLATING TERM)

Listed in CP-violation section below.

QUADRATIC COEFFICIENT  $h$  FOR  $K_L^0 \rightarrow \pi^0 \pi^0 \pi^0$ 

We do not average measurements that do not account for the effect of final state rescattering.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN
<b><math>+0.59 \pm 0.20 \pm 1.16</math></b>	68M	<sup>1</sup> ABOUZAID	08A KTEV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-6.1 \pm 0.9 \pm 0.5$	14.7M	<sup>2</sup> LAI	01B NA48
$-3.3 \pm 1.1 \pm 0.7$	5M	<sup>2,3</sup> SOMALWAR	92 E731

- <sup>1</sup> Result obtained using C13pi model of CABIBBO 05 to include  $\pi\pi$  rescattering effects. The systematic error includes an external error of  $1.06 \times 10^{-3}$  from the parametrization input of  $(a_0 - a_2) m_{\pi^+} = 0.268 \pm 0.017$  from BATLEY 06B.
- <sup>2</sup> LAI 01B and SOMALWAR 92 results do not include  $\pi\pi$  final state rescattering effects.
- <sup>3</sup> SOMALWAR 92 chose  $m_{\pi^+}$  as normalization to make it compatible with the Particle Data Group  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$  definitions.

# Meson Particle Listings

$K_L^0$

## $K_L^0$ FORM FACTORS

For discussion, see note on form factors in the  $K^\pm$  section of the Particle Listings above.

In the form factor comments, the following symbols are used.

- $f_+$  and  $f_-$  are form factors for the vector matrix element.
- $f_S$  and  $f_T$  refer to the scalar and tensor terms.
- $f_0(t) = f_+(t) + f_-(t) t/(m_{K^0}^2 - m_{\pi^+}^2)$ .
- $t$  = momentum transfer to the  $\pi$ .
- $\lambda_+$  and  $\lambda_0$  are the linear expansion coefficients of  $f_+$  and  $f_0$ :
- $f_+(t) = f_+(0) (1 + \lambda_+ t/m_{\pi^+}^2)$
- For quadratic expansion
- $f_+(t) = f_+(0) (1 + \lambda'_+ t/m_{\pi^+}^2 + \frac{\lambda''_+}{2} t^2/m_{\pi^+}^4)$
- as used by KTeV. If there is a non-vanishing quadratic term, then  $\lambda_+$  represents an average slope, which is then different from  $\lambda'_+$ .
- NA48 ( $K_{e3}$ ) and ISTRA quadratic expansion coefficients are converted with  $\lambda'_+{}^{PDG} = \lambda_+{}^{NA48}$  and  $\lambda''_+{}^{PDG} = 2 \lambda'_+{}^{NA48}$
- $\lambda'_+{}^{PDG} = (\frac{m_{\pi^+}}{m_{\pi^0}})^2 \lambda_+{}^{ISTRA}$  and
- $\lambda''_+{}^{PDG} = 2 (\frac{m_{\pi^+}}{m_{\pi^0}})^4 \lambda'_+{}^{ISTRA}$
- ISTRA linear expansion coefficients are converted with  $\lambda_+{}^{PDG} = (\frac{m_{\pi^+}}{m_{\pi^0}})^2 \lambda_+{}^{ISTRA}$  and  $\lambda_0{}^{PDG} = (\frac{m_{\pi^+}}{m_{\pi^0}})^2 \lambda_0{}^{ISTRA}$

The pole parametrization is

$$f_+(t) = f_+(0) \left( \frac{M_V^2}{M_V^2 - t} \right)$$

$$f_0(t) = f_0(0) \left( \frac{M_S^2}{M_S^2 - t} \right)$$

where  $M_V$  and  $M_S$  are the vector and scalar pole masses.

The dispersive parametrization is

$$f_+(t) = f_+(0) \exp \left[ \frac{t}{m_\pi^2} (\Lambda_+ + H(t)) \right];$$

$$f_0(t) = f_+(0) \exp \left[ \frac{t}{m_K^2 - m_\pi^2} (\ln[C] - G(t)) \right],$$

where  $\Lambda_+$  is the slope parameter and  $\ln[C] = \ln[f_0(m_K^2 - m_\pi^2)]$  is the logarithm of the scalar form factor at the Callan-Treiman point.

$H(t)$  and  $G(t)$  are dispersive integrals.

The following abbreviations are used:

DP = Dalitz plot analysis.

PI =  $\pi$  spectrum analysis.

MU =  $\mu$  spectrum analysis.

POL =  $\mu$  polarization analysis.

BR =  $K_{\mu 3}^0/K_{e 3}^0$  branching ratio analysis.

E = positron or electron spectrum analysis.

RC = radiative corrections.

## $\lambda_+$ (LINEAR ENERGY DEPENDENCE OF $f_+$ IN $K_{e 3}^0$ DECAY)

For radiative correction of  $K_{e 3}^0$  DP, see GINSBERG 67, BECHERRAWY 70, CIRIGLIANO 02, CIRIGLIANO 04, and ANDRE 07. Results labeled OUR FIT are discussed in the review “ $K_{e 3}^\pm$  and  $K_{\mu 3}^0$  Form Factors” in the  $K^\pm$  Listings. For earlier, lower statistics results, see the 2004 edition of this review, Physics Letters **B592** 1 (2004).

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.82 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1. Assuming $\mu$ -e universality
<b>2.85 ± 0.04 OUR AVERAGE</b>				
2.86 ± 0.05 ± 0.04	2M	AMBROSINO 06D	KLOE	
2.832 ± 0.037 ± 0.043	1.9M	ALEXOPOU... 04A	KTEV	PI, no $\mu = e$
2.88 ± 0.04 ± 0.11	5.6M	<sup>1</sup> LAI 04C	NA48	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.84 ± 0.07 ± 0.13	5.6M	<sup>2</sup> LAI 04C	NA48	DP
2.45 ± 0.12 ± 0.22	366k	APOSTOLA... 00	CPLR	DP
3.06 ± 0.34	74k	BIRULEV 81	SPEC	DP
3.12 ± 0.25	500k	GJESDAL 76	SPEC	DP
2.70 ± 0.28	25k	BLUMENTHAL 75	SPEC	DP

<sup>1</sup> Results from linear fit and assuming only vector and axial couplings.

<sup>2</sup> Results from linear fit with  $|f_S/f_+|$  and  $|f_T/f_+|$  free.

## $\lambda_+$ (LINEAR ENERGY DEPENDENCE OF $f_+$ IN $K_{\mu 3}^0$ DECAY)

Results labeled OUR FIT are discussed in the review “ $K_{e 3}^\pm$  and  $K_{\mu 3}^0$  Form Factors” in the  $K^\pm$  Listings. For earlier, lower statistics results, see the 2004 edition of this review, Physics Letters **B592** 1 (2004).

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.82 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1. Assuming $\mu$ -e universality
<b>2.71 ± 0.10 OUR FIT</b>				Error includes scale factor of 1.4. Not assuming $\mu$ -e universality
2.67 ± 0.06 ± 0.08	2.3M	<sup>1</sup> LAI 07A	NA48	DP
2.745 ± 0.088 ± 0.063	1.5M	ALEXOPOU... 04A	KTEV	DP, no $\mu = e$
2.813 ± 0.051	3.4M	ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$
3.0 ± 0.3	1.6M	DONALDSON 74B	SPEC	DP
• • • We do not use the following data for averages, fits, limits, etc. • • •				

4.27 ± 0.44 150k BIRULEV 81 SPEC DP

<sup>1</sup> LAI 07A gives a correlation  $-0.40$  between their  $\lambda_0$  and  $\lambda_+$  measurements.

## $\lambda_0$ (LINEAR ENERGY DEPENDENCE OF $f_0$ IN $K_{\mu 3}^0$ DECAY)

Wherever possible, we have converted the above values of  $\xi(0)$  into values of  $\lambda_0$  using the associated  $\lambda_+^\mu$  and  $d\xi(0)/d\lambda_+$ . Results labeled OUR FIT are discussed in the review “ $K_{e 3}^\pm$  and  $K_{\mu 3}^0$  Form Factors” in the  $K^\pm$  Listings. For earlier, lower statistics results, see the 2004 edition of this review, Physics Letters **B592** 1 (2004).

VALUE (units $10^{-2}$ )	$d\lambda_0/d\lambda_+$	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.38 ± 0.18 OUR FIT</b>					Error includes scale factor of 2.2. Assuming $\mu$ -e universality
<b>1.42 ± 0.23 OUR FIT</b>					Error includes scale factor of 2.8. Not assuming $\mu$ -e universality
1.17 ± 0.07 ± 0.10		2.3M	<sup>1</sup> LAI 07A	NA48	DP
1.657 ± 0.125	$-0.44$	1.5M	<sup>2</sup> ALEXOPOU... 04A	KTEV	DP, no $\mu = e$
1.635 ± 0.121	$-0.85$	3.4M	<sup>3</sup> ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$
+1.9 ± 0.4	$-0.47$	1.6M	<sup>4</sup> DONALDSON 74B	SPEC	DP

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.41 ± 0.67 unknown 150k <sup>5</sup> BIRULEV 81 SPEC DP

<sup>1</sup> LAI 07A gives a correlation  $-0.40$  between their  $\lambda_0$  and  $\lambda_+$  measurements.

<sup>2</sup> ALEXOPOULOS 04A gives a correlation  $-0.38$  between their  $\lambda_0$  and  $\lambda_+$  measurements.

<sup>3</sup> ALEXOPOULOS 04A gives a correlation  $-0.36$  between their  $\lambda_0$  and  $\lambda_+$  measurements.

<sup>4</sup> DONALDSON 74B  $d\lambda_0/d\lambda_+$  obtained from figure 18.

<sup>5</sup> BIRULEV 81 gives  $d\lambda_0/d\lambda_+ = -1.5$ , giving an unreasonably narrow error ellipse which dominates all other results. We use  $d\lambda_0/d\lambda_+ = 0$ .

## $\lambda'_+ ($ LINEAR $K_{e 3}^0$ FORM FACTOR FROM QUADRATIC FIT)

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.40 ± 0.12 OUR FIT</b>				Error includes scale factor of 1.2. Assuming $\mu$ -e universality
<b>2.49 ± 0.13 OUR FIT</b>				Error includes scale factor of 1.1. Not assuming $\mu$ -e universality
<b>2.48 ± 0.17 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
2.55 ± 0.15 ± 0.10	2M	<sup>1</sup> AMBROSINO 06D	KLOE	
2.167 ± 0.137 ± 0.143	1.9M	<sup>2</sup> ALEXOPOU... 04A	KTEV	PI, no $\mu = e$
2.80 ± 0.19 ± 0.15	5.6M	<sup>3</sup> LAI 04C	NA48	DP

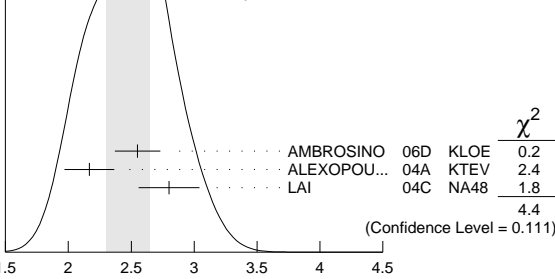
<sup>1</sup> We use AMBROSINO 06D result in the fit not assuming  $\mu$ -e universality. This result enters the fit assuming  $\mu$ -e universality via AMBROSINO 07C measurement of  $\lambda'_+$  in  $K_{\mu 3}$  decays. AMBROSINO 06D gives a correlation  $-0.95$  between their  $\lambda'_+$  and  $\lambda''_+$ .

<sup>2</sup> ALEXOPOULOS 04A gives a correlation  $-0.97$  between their  $\lambda'_+$  and  $\lambda''_+$ .

<sup>3</sup> For LAI 04C we calculate a correlation  $-0.88$  between their  $\lambda'_+$  and  $\lambda''_+$ .

WEIGHTED AVERAGE  
2.48±0.17 (Error scaled by 1.5)

Values above of weighted average, error, and scale factor are based upon the data in this ideogram only. They are not necessarily the same as our ‘best’ values, obtained from a least-squares constrained fit utilizing measurements of other (related) quantities as additional information.



## $\lambda''_+ ($ QUADRATIC $K_{e 3}^0$ FORM FACTOR)

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.20 ± 0.05 OUR FIT</b>				Error includes scale factor of 1.2. Assuming $\mu$ -e universality
<b>0.16 ± 0.05 OUR FIT</b>				Error includes scale factor of 1.1. Not assuming $\mu$ -e universality
<b>0.17 ± 0.07 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
0.14 ± 0.07 ± 0.04	2M	<sup>1</sup> AMBROSINO 06D	KLOE	
0.287 ± 0.057 ± 0.053	1.9M	<sup>2</sup> ALEXOPOU... 04A	KTEV	PI, no $\mu = e$
0.04 ± 0.08 ± 0.04	5.6M	<sup>3,4</sup> LAI 04C	NA48	DP

<sup>1</sup> We use AMBROSINO 06D result in the fit not assuming  $\mu$ -e universality. This result enters the fit assuming  $\mu$ -e universality via AMBROSINO 07C measurement of  $\lambda'_+$  in  $K_{\mu 3}$  decays. AMBROSINO 06D gives a correlation  $-0.95$  between their  $\lambda'_+$  and  $\lambda''_+$ .

<sup>2</sup> ALEXOPOULOS 04A gives a correlation  $-0.97$  between their  $\lambda'_+$  and  $\lambda''_+$ .

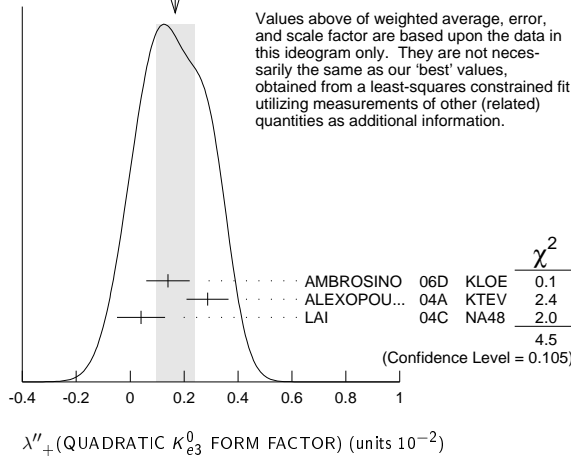
<sup>3</sup> Values doubled to agree with PDG conventions described above.

<sup>4</sup> LAI 04C gives a correlation  $-0.88$  between their  $\lambda'_+$  and  $\lambda''_+$ .



See key on page 601

## Meson Particle Listings

 $K_L^0$ WEIGHTED AVERAGE  
0.17±0.07 (Error scaled by 1.5) $\lambda'_+$  (LINEAR  $K_{\mu 3}^0$  FORM FACTOR FROM QUADRATIC FIT)

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.40 ± 0.12 OUR FIT</b>				Error includes scale factor of 1.2. Assuming $\mu$ -e universality
<b>1.89 ± 0.24 OUR FIT</b>				Not assuming $\mu$ -e universality
2.23 ± 0.98 ± 0.37	1.8M	<sup>1</sup> AMBROSINO 07C	KLOE	no $\mu = e$
2.56 ± 0.15 ± 0.09	3.8M	<sup>1</sup> AMBROSINO 07C	KLOE	$\mu = e$
2.05 ± 0.22 ± 0.24	2.3M	<sup>1</sup> LAI 07A	NA48	DP
1.703 ± 0.319 ± 0.177	1.5M	<sup>1</sup> ALEXOPOU... 04A	KTEV	DP, no $\mu = e$
2.064 ± 0.175	3.4M	<sup>1</sup> ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$

<sup>1</sup> See section  $\lambda_0$  below for correlations. $\lambda''_+$  (QUADRATIC  $K_{\mu 3}^0$  FORM FACTOR)

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.20 ± 0.05 OUR FIT</b>				Error includes scale factor of 1.2. Assuming $\mu$ -e universality
<b>0.37 ± 0.12 OUR FIT</b>				Error includes scale factor of 1.3. Not assuming $\mu$ -e universality
0.48 ± 0.49 ± 0.16	1.8M	<sup>1</sup> AMBROSINO 07C	KLOE	no $\mu = e$
0.15 ± 0.07 ± 0.04	3.8M	<sup>1</sup> AMBROSINO 07C	KLOE	$\mu = e$
0.26 ± 0.09 ± 0.10	2.3M	<sup>1</sup> LAI 07A	NA48	DP
0.443 ± 0.131 ± 0.072	1.5M	<sup>1</sup> ALEXOPOU... 04A	KTEV	DP, no $\mu = e$
0.320 ± 0.069	3.4M	<sup>1</sup> ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$

<sup>1</sup> See section  $\lambda_0$  below for correlations. $\lambda_0$  (LINEAR  $f_0 K_{\mu 3}^0$  FORM FACTOR FROM QUADRATIC FIT)

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.16 ± 0.09 OUR FIT</b>				Error includes scale factor of 1.2. Assuming $\mu$ -e universality
<b>1.07 ± 0.14 OUR FIT</b>				Error includes scale factor of 1.3. Not assuming $\mu$ -e universality
0.91 ± 0.59 ± 0.26	1.8M	<sup>1</sup> AMBROSINO 07C	KLOE	no $\mu = e$
1.54 ± 0.18 ± 0.13	3.8M	<sup>2</sup> AMBROSINO 07C	KLOE	$\mu = e$
0.95 ± 0.11 ± 0.08	2.3M	<sup>3</sup> LAI 07A	NA48	DP
1.281 ± 0.136 ± 0.122	1.5M	<sup>4</sup> ALEXOPOU... 04A	KTEV	DP, no $\mu = e$
1.372 ± 0.131	3.4M	<sup>5</sup> ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$

<sup>1</sup> AMBROSINO 07C, not assuming  $\mu$ -e universality, gives a correlation matrix

$$\begin{array}{cc} \lambda'_+ & \lambda''_+ \\ \lambda'_+ & -0.97 & 1 \\ \lambda_0 & 0.81 & -0.91 \end{array}$$

<sup>2</sup> AMBROSINO 07C, assuming  $\mu$ -e universality, gives a correlation matrix

$$\begin{array}{cc} \lambda'_+ & \lambda''_+ \\ \lambda'_+ & -0.95 & 1 \\ \lambda_0 & 0.29 & -0.38 \end{array}$$

<sup>3</sup> LAI 07A gives a correlation matrix

$$\begin{array}{cc} \lambda'_+ & \lambda''_+ \\ \lambda'_+ & -0.96 & 1 \\ \lambda_0 & 0.63 & -0.73 \end{array}$$

<sup>4</sup> ALEXOPOULOS 04A, not assuming  $\mu$ -e universality, gives a correlation matrix

$$\begin{array}{ccc} \lambda'_+ & \lambda''_+ & \lambda_0 \\ \lambda'_+ & 1 & \\ \lambda''_+ & -0.96 & 1 \\ \lambda_0 & 0.65 & -0.75 & 1 \end{array}$$

<sup>5</sup> ALEXOPOULOS 04A, assuming  $\mu$ -e universality, gives a correlation matrix

$$\begin{array}{ccc} \lambda'_+ & \lambda''_+ & \lambda_0 \\ \lambda'_+ & 1 & \\ \lambda''_+ & -0.97 & 1 \\ \lambda_0 & 0.34 & -0.44 & 1 \end{array}$$

 $M_V^e$  (POLE MASS FOR  $K_{e 3}^0$  DECAY)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>878 ± 6 OUR FIT</b>				Error includes scale factor of 1.1. Assuming $\mu$ -e universality
<b>875 ± 5 OUR AVERAGE</b>				
870 ± 6 ± 7	2M	AMBROSINO	06D	KLOE
881.03 ± 5.12 ± 4.94	1.9M	ALEXOPOU...	04A	KTEV PI, no $\mu = e$
859 ± 18	5.6M	LAI	04C	NA48

 $M_V^\mu$  (POLE MASS FOR  $K_{\mu 3}^0$  DECAY)

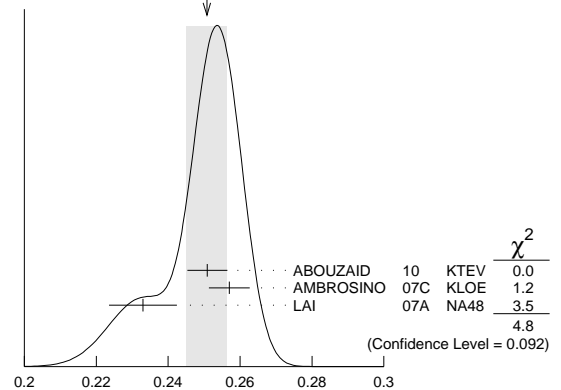
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>878 ± 6 OUR FIT</b>				Error includes scale factor of 1.1. Assuming $\mu$ -e universality
<b>900 ± 21 OUR FIT</b>				Error includes scale factor of 1.7. Not assuming $\mu$ -e universality
905 ± 9 ± 17	2.3M	<sup>1</sup> LAI	07A	NA48 DP
889.19 ± 12.81 ± 9.92	1.5M	<sup>1</sup> ALEXOPOU... 04A	KTEV	DP, no $\mu = e$
882.32 ± 6.54	3.4M	<sup>1</sup> ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$

<sup>1</sup> See section  $M_S^\mu$  below for correlations. $M_S^\mu$  (POLE MASS FOR  $K_{\mu 3}^0$  DECAY)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1252 ± 90 OUR FIT</b>				Error includes scale factor of 2.6. Assuming $\mu$ -e universality
<b>1222 ± 80 OUR FIT</b>				Error includes scale factor of 2.3. Not assuming $\mu$ -e universality
1400 ± 46 ± 53	2.3M	<sup>1</sup> LAI	07A	NA48 DP
1167.14 ± 28.30 ± 31.04	1.5M	<sup>2</sup> ALEXOPOU... 04A	KTEV	PI, no $\mu = e$
1173.80 ± 39.47	3.4M	<sup>3</sup> ALEXOPOU... 04A	KTEV	PI, DP, $\mu = e$

<sup>1</sup> LAI 07A gives a correlation  $-0.47$  between their  $M_S^\mu$  and  $M_V^\mu$  measurements, not assuming  $\mu$ -e universality.<sup>2</sup> ALEXOPOULOS 04A gives a correlation  $-0.46$  between their  $M_S^\mu$  and  $M_V^\mu$  and measurements, not assuming  $\mu$ -e universality.<sup>3</sup> ALEXOPOULOS 04A gives a correlation  $-0.40$  between their  $M_S^\mu$  and  $M_V^\mu$  and measurements, assuming  $\mu$ -e universality. $\Lambda_+$  (DISPERSIVE VECTOR FORM FACTOR FOR  $K_{\mu 3}^0$  DECAY)See the review on " $K_{\ell 3}^\pm$  and  $K_{\ell 3}^0$  Form Factors" for details of the dispersive parametrization.

VALUE (units $10^{-1}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.251 ± 0.006 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
0.2509 ± 0.0035 ± 0.0043	3.4M	<sup>1</sup> ABOUZAID 10	KTEV	$\mu = e$
0.257 ± 0.004 ± 0.004	3.8M	<sup>2</sup> AMBROSINO 07C	KLOE	$\mu = e$
0.233 ± 0.005 ± 0.008	2.3M	<sup>3</sup> LAI 07A	NA48	DP

<sup>1</sup> Obtained from a sample of 1.9 M  $K_{e 3}$  and 1.5 M  $K_{\mu 3}$ . The correlation between  $\Lambda_+$  and  $\ln(C)$  is  $-0.269$ .<sup>2</sup> AMBROSINO 07C results include 2M  $K_{e 3}$  events from AMBROSINO 06D. The correlation between  $\Lambda_+$  and  $\ln(C)$  is  $-0.26$ .<sup>3</sup> LAI 07A gives a correlation  $-0.44$  between their  $\Lambda_+$  and  $\ln(C)$  measurements.WEIGHTED AVERAGE  
0.251±0.006 (Error scaled by 1.5) $\ln(C)$  (DISPERSIVE SCALAR FORM FACTOR FOR  $K_{\mu 3}^0$  DECAY)See the review on " $K_{\ell 3}^\pm$  and  $K_{\ell 3}^0$  Form Factors" for details of the dispersive parametrization.

VALUE (units $10^{-1}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.75 ± 0.18 OUR AVERAGE</b>				Error includes scale factor of 2.0. See the ideogram below.
1.915 ± 0.078 ± 0.094	3.4M	<sup>1</sup> ABOUZAID 10	KTEV	$\mu = e$
2.04 ± 0.19 ± 0.15	3.8M	<sup>2</sup> AMBROSINO 07C	KLOE	$\mu = e$
1.438 ± 0.080 ± 0.112	2.3M	<sup>3</sup> LAI	07A	NA48 DP

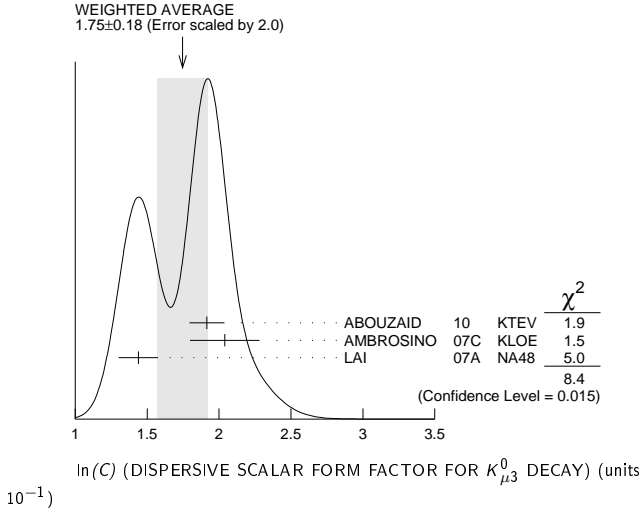
# Meson Particle Listings

$K_L^0$

<sup>1</sup> Obtained from a sample of 1.9 M  $K_{e3}$  and 1.5 M  $K_{\mu 3}$ . The correlation between  $\Lambda_+$  and  $\ln(C)$  is  $-0.269$ .

<sup>2</sup> AMBROSINO 07C results include 2M  $K_{e3}$  events from AMBROSINO 06D. We convert  $(\Lambda_+, \Lambda_0)$  to  $(\Lambda_+, \ln(C))$  parametrization using  $\ln(C) = (\Lambda_0 \cdot 11.713 + 0.0398) \pm 0.0041$ , where the error is due to theory parametrization of the form factor. The correlation between  $\Lambda_+$  and  $\ln(C)$  is  $-0.26$ .

<sup>3</sup> LAI 07A gives a correlation  $-0.44$  between their  $\Lambda_+$  and  $\ln(C)$  measurements.



## $a_1(t_0, Q^2)$ FORM FACTOR PARAMETER

See HILL 06 for a definition of this parameter.

VALUE	EVTS	DOCUMENT ID	TECN
<b>1.023 ± 0.028 ± 0.029</b>	2M	<sup>1</sup> ABOUZAID 06c	KTEV

<sup>1</sup>  $Q^2 = 2 \text{ GeV}^2$ ,  $t_0 = 0.49 (m_K - m_\pi)^2$ . Correlation between  $a_1$  and  $a_2$ :  $\rho_{12} = -0.064$ .

## $a_2(t_0, Q^2)$ FORM FACTOR PARAMETER

See HILL 06 for a definition of this parameter.

VALUE	EVTS	DOCUMENT ID	TECN
<b>0.75 ± 1.58 ± 1.47</b>	2M	<sup>1</sup> ABOUZAID 06c	KTEV

<sup>1</sup>  $Q^2 = 2 \text{ GeV}^2$ ,  $t_0 = 0.49 (m_K - m_\pi)^2$ . Correlation between  $a_1$  and  $a_2$ :  $\rho_{12} = -0.064$ .

## $|f_S/f_+|$ FOR $K_{e3}^0$ DECAY

Ratio of scalar to  $f_+$  couplings.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.5 ± 0.7 ± 1.2</b>	5.6M	<sup>1</sup> LAI	04c	NA48	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<9.5	95	18k	HILL	78	STRC	See also BIRULEV 81
<7.	68	48k	BIRULEV	76	SPEC	
<4.	68	25k	BLUMENTHAL75	SPEC		

<sup>1</sup> Results from linear fit with  $|f_S/f_+|$  and  $|f_T/f_+|$  free.

## $|f_T/f_+|$ FOR $K_{e3}^0$ DECAY

Ratio of tensor to  $f_+$  couplings.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5 ± 3 ± 3</b>	5.6M	<sup>1</sup> LAI	04c	NA48	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<40.	95	18k	HILL	78	STRC	See also BIRULEV 81
<34.	68	48k	BIRULEV	76	SPEC	
<23.	68	25k	BLUMENTHAL75	SPEC		

<sup>1</sup> Results from linear fit with  $|f_S/f_+|$  and  $|f_T/f_+|$  free.

## $|f_T/f_+|$ FOR $K_{\mu 3}^0$ DECAY

Ratio of tensor to  $f_+$  couplings.

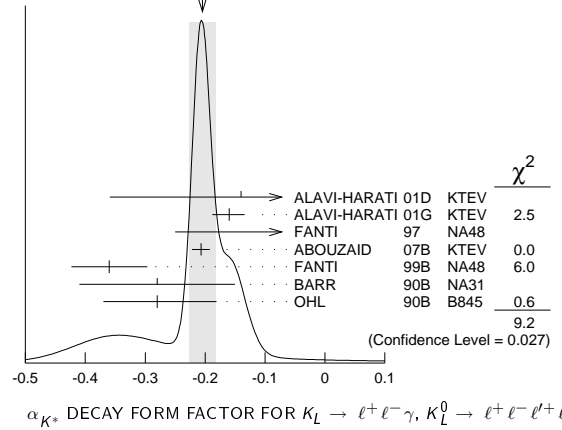
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN
<b>12 ± 12.</b>	BIRULEV 81	SPEC

## $\alpha_{K^*}$ DECAY FORM FACTOR FOR $K_L \rightarrow \ell^+ \ell^- \gamma$ , $K_L^0 \rightarrow \ell^+ \ell^- \ell'^+ \ell'^-$

Average of all  $\alpha_{K^*}$  measurements (from each of three datablocks following this one) assuming lepton universality.

VALUE	DOCUMENT ID
<b>-0.205 ± 0.022 OUR AVERAGE</b>	Includes data from the 3 datablocks that follow this one. Error includes scale factor of 1.8. See the ideogram below.

WEIGHTED AVERAGE  
-0.205±0.022 (Error scaled by 1.8)



## $\alpha_{K^*}$ DECAY FORM FACTOR FOR $K_L \rightarrow e^+ e^- \gamma$

$\alpha_{K^*}$  is the constant in the model of BERGSTROM 83 which measures the relative strength of the vector-vector transition  $K_L \rightarrow K^* \gamma$  with  $K^* \rightarrow \rho, \omega, \phi \rightarrow \gamma^*$  and the pseudoscalar-pseudoscalar transition  $K_L \rightarrow \pi, \eta, \eta' \rightarrow \gamma \gamma^*$ .

VALUE	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

**-0.217 ± 0.034 OUR AVERAGE** Error includes scale factor of 2.4.

-0.207 ± 0.012 ± 0.009	83k	<sup>1</sup> ABOUZAID 07b	KTEV
-0.36 ± 0.06 ± 0.02	6864	FANTI 99b	NA48
-0.28 ± 0.13		BARR 90b	NA31
-0.280 + 0.099 - 0.090		OHL 90b	B845

<sup>1</sup> ABOUZAID 07b measures  $C$ .  $\alpha_{K^*} = -0.517 \pm 0.030 \pm 0.022$ . We assume  $C = 2.5$ , as in all other measurements.

## $\alpha_{K^*}$ DECAY FORM FACTOR FOR $K_L \rightarrow \mu^+ \mu^- \gamma$

$\alpha_{K^*}$  is the constant in the model of BERGSTROM 83 described in the previous section.

VALUE	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

**-0.158 ± 0.027 OUR AVERAGE**

-0.160 + 0.026 - 0.028	9100	ALAVI-HARATI 01G	KTEV
-0.04 + 0.24 - 0.21		FANTI 97	NA48

## $\alpha_{K^*}^{\text{eff}}$ DECAY FORM FACTOR FOR $K_L \rightarrow e^+ e^- e^+ e^-$

$\alpha_{K^*}^{\text{eff}}$  is the parameter describing the relative strength of an intermediate pseudoscalar decay amplitude and a vector meson decay amplitude in the model of BERGSTROM 83. It takes into account both the radiative effects and the form factor. Since there are two  $e^+ e^-$  pairs here compared with one in  $e^+ e^- \gamma$  decays, a factorized expression is used for the  $e^+ e^- e^+ e^-$  decay form factor.

VALUE	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

**-0.14 ± 0.16 ± 0.15** 441 ALAVI-HARATI 01D KTEV

## $\alpha_{DIP}$ DECAY FORM FACTOR FOR $K_L^0 \rightarrow \ell^+ \ell^- \gamma$ , $K_L^0 \rightarrow \ell^+ \ell^- \ell'^+ \ell'^-$

Average of all  $\alpha_{DIP}$  measurements (from each of three datablocks following this one) assuming lepton universality.

VALUE	DOCUMENT ID
<b>-1.69 ± 0.08 OUR AVERAGE</b>	Includes data from the 3 datablocks that follow this one. Error includes scale factor of 1.7.

## $\alpha_{DIP}$ DECAY FORM FACTOR FOR $K_L^0 \rightarrow e^+ e^- \gamma$

$\alpha_{DIP}$  parameter in  $K_L^0 \rightarrow \gamma^* \gamma^*$  form factor by DAMBROSIO 98, motivated by vector meson dominance and a proper short distance behavior.

VALUE	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

**-1.729 ± 0.043 ± 0.028** 83k ABOUZAID 07b KTEV

## $\alpha_{DIP}$ DECAY FORM FACTOR FOR $K_L^0 \rightarrow \mu^+ \mu^- \gamma$

$\alpha_{DIP}$  is a constant in the model of DAMBROSIO 98 described in the previous section.

VALUE	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

**-1.54 ± 0.10** 9100 ALAVI-HARATI 01G KTEV

## $\alpha_{DIP}$ DECAY FORM FACTOR FOR $K_L^0 \rightarrow e^+ e^- \mu^+ \mu^-$

$\alpha_{DIP}$  is a constant in the model of DAMBROSIO 98 described in the previous section.

VALUE	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

**-1.59 ± 0.37** 131 ALAVI-HARATI 03B KTEV

**$a_1/a_2$  FORM FACTOR FOR  $M1$  DIRECT EMISSION AMPLITUDE**Form factor =  $\tilde{g}_{M1} \left[ 1 + \frac{a_1/a_2}{(M_\rho^2 - M_K^2) + 2M_K E_\gamma} \right]$  as described in ALAVI-HARATI 00b.

VALUE (GeV <sup>2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.737 ± 0.014 OUR AVERAGE</b>				
-0.744 ± 0.027 ± 0.032	5241	1 ABOUZAID	06	KTEV $\pi^+ \pi^- e^+ e^-$
-0.738 ± 0.007 ± 0.018	111k	2 ABOUZAID	06A	KTEV $\pi^+ \pi^+ \gamma$
-0.81 ± 0.07 ± 0.02		3 LAI	03c	NA48 $\pi^+ \pi^- e^+ e^-$
-0.737 ± 0.026 ± 0.022		4 ALAVI-HARATI 01B		$\pi^+ \pi^- \gamma$
-0.720 ± 0.028 ± 0.009	1766	5 ALAVI-HARATI 00b	KTEV	$\pi^+ \pi^- e^+ e^-$

<sup>1</sup> ABOUZAID 06 also measured  $|\tilde{g}_{M1}| = 1.11 \pm 0.14$ .<sup>2</sup> ABOUZAID 06a also measured  $|\tilde{g}_{M1}| = 1.198 \pm 0.035 \pm 0.086$ .<sup>3</sup> LAI 03c also measured  $\tilde{g}_{M1} = 0.99 \pm 0.28 \pm 0.07$ .<sup>4</sup> ALAVI-HARATI 01B fit gives  $\chi^2/\text{DOF} = 38.8/27$ . Linear and quadratic fits give  $\chi^2/\text{DOF} = 43.2/27$  and  $37.6/26$  respectively.<sup>5</sup> ALAVI-HARATI 00b also measured  $|\tilde{g}_{M1}| = 1.35 \pm 0.20 \pm 0.17$ . **$\mathcal{T}_S$  DECAY FORM FACTOR FOR  $K_L^0 \rightarrow \pi^\pm \pi^0 e^\mp \nu_e$** 

VALUE	DOCUMENT ID	TECN
<b>0.049 ± 0.011 OUR AVERAGE</b>	Error includes scale factor of 1.7.	
0.052 ± 0.006 ± 0.002	BATLEY	04 NA48
0.010 ± 0.016 ± 0.017	MAKOFF	93 E731

 **$\mathcal{T}_P$  DECAY FORM FACTOR FOR  $K_L^0 \rightarrow \pi^\pm \pi^0 e^\mp \nu_e$** 

VALUE	DOCUMENT ID	TECN
<b>-0.052 ± 0.012 OUR AVERAGE</b>		
-0.051 ± 0.011 ± 0.005	BATLEY	04 NA48
-0.079 ± 0.049 ± 0.022	MAKOFF	93 E731

 **$\lambda_g$  DECAY FORM FACTOR FOR  $K_L^0 \rightarrow \pi^\pm \pi^0 e^\mp \nu_e$** 

VALUE	DOCUMENT ID	TECN
<b>0.085 ± 0.020 OUR AVERAGE</b>		
0.087 ± 0.019 ± 0.006	BATLEY	04 NA48
0.014 ± 0.087 ± 0.070	MAKOFF	93 E731

 **$\tilde{h}$  DECAY FORM FACTOR FOR  $K_L^0 \rightarrow \pi^\pm \pi^0 e^\mp \nu_e$** 

VALUE	DOCUMENT ID	TECN
<b>-0.30 ± 0.13 OUR AVERAGE</b>		
-0.32 ± 0.12 ± 0.07	BATLEY	04 NA48
-0.07 ± 0.31 ± 0.31	MAKOFF	93 E731

 **$L_3$  CHIRAL PERT. THEO. PARAM. FOR  $K_L^0 \rightarrow \pi^\pm \pi^0 e^\mp \nu_e$** 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN
<b>-3.96 ± 0.28 OUR AVERAGE</b>	Error includes scale factor of 1.6.	
-4.1 ± 0.2	BATLEY	04 NA48
-3.4 ± 0.4	1 MAKOFF	93 E731

<sup>1</sup> MAKOFF 93 sign has been changed to negative to agree with the sign convention used in BATLEY 04. **$a_V$ , VECTOR MESON EXCHANGE CONTRIBUTION**

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.43 ± 0.06 OUR AVERAGE</b>	Error includes scale factor of 1.5.			
-0.31 ± 0.05 ± 0.07	1.4k	1 ABOUZAID	08	KTEV
-0.46 ± 0.03 ± 0.04		LAI	02B	NA48 $K_L^0 \rightarrow \pi^0 2\gamma$
-0.67 ± 0.21 ± 0.12		ALAVI-HARATI 01E	KTEV	$K_L^0 \rightarrow \pi^0 e^+ e^- \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.72 ± 0.05 ± 0.06		2 ALAVI-HARATI 99B	KTEV	$K_L^0 \rightarrow \pi^0 2\gamma$

<sup>1</sup> Using KTeV dataset collected in 1996, 1997, and 1999.<sup>2</sup> Superseded by ABOUZAID 08. **$CP$  VIOLATION IN  $K_L$  DECAYS**

Updated April 2016 by L. Wolfenstein (Carnegie-Mellon University), C.-J. Lin (LBNL), and T.G. Trippe (LBNL).

The symmetries  $C$  (particle-antiparticle interchange) and  $P$  (space inversion) hold for strong and electromagnetic interactions. After the discovery of large  $C$  and  $P$  violation in the weak interactions, it appeared that the product  $CP$  was a good symmetry. In 1964  $CP$  violation was observed in  $K^0$  decays at a level given by the parameter  $\epsilon \approx 2.3 \times 10^{-3}$ .

A unified treatment of  $CP$  violation in  $K$ ,  $D$ ,  $B$ , and  $B_s$  mesons is given in “ $CP$  Violation in Meson Decays” by D. Kirkby and Y. Nir in this *Review*. A more detailed review including a thorough discussion of the experimental techniques used to determine  $CP$  violation parameters is given in a book

by K. Kleinknecht [1]. Here we give a concise summary of the formalism needed to define the parameters of  $CP$  violation in  $K_L$  decays, and a description of our fits for the best values of these parameters.

**1. Formalism for  $CP$  violation in Kaon decay:**

$CP$  violation has been observed in the semi-leptonic decays  $K_L^0 \rightarrow \pi^\mp \ell^\pm \nu$ , and in the nonleptonic decay  $K_L^0 \rightarrow 2\pi$ . The experimental numbers that have been measured are

$$A_L = \frac{\Gamma(K_L^0 \rightarrow \pi^- \ell^+ \nu) - \Gamma(K_L^0 \rightarrow \pi^+ \ell^- \nu)}{\Gamma(K_L^0 \rightarrow \pi^- \ell^+ \nu) + \Gamma(K_L^0 \rightarrow \pi^+ \ell^- \nu)} \quad (1a)$$

$$\eta_{+-} = A(K_L^0 \rightarrow \pi^+ \pi^-) / A(K_S^0 \rightarrow \pi^+ \pi^-) = |\eta_{+-}| e^{i\phi_{+-}} \quad (1b)$$

$$\eta_{00} = A(K_L^0 \rightarrow \pi^0 \pi^0) / A(K_S^0 \rightarrow \pi^0 \pi^0) = |\eta_{00}| e^{i\phi_{00}} \quad (1c)$$

$CP$  violation can occur either in the  $K^0 - \bar{K}^0$  mixing or in the decay amplitudes. Assuming  $CPT$  invariance, the mass eigenstates of the  $K^0 - \bar{K}^0$  system can be written

$$|K_S\rangle = p|K^0\rangle + q|\bar{K}^0\rangle, \quad |K_L\rangle = p|K^0\rangle - q|\bar{K}^0\rangle \quad (2)$$

If  $CP$  invariance held, we would have  $q = p$  so that  $K_S$  would be  $CP$ -even and  $K_L$   $CP$ -odd. (We define  $|\bar{K}^0\rangle$  as  $CP$   $|K^0\rangle$ ).  $CP$  violation in  $K^0 - \bar{K}^0$  mixing is then given by the parameter  $\tilde{\epsilon}$  where

$$\frac{p}{q} = \frac{(1 + \tilde{\epsilon})}{(1 - \tilde{\epsilon})} \quad (3)$$

$CP$  violation can also occur in the decay amplitudes

$$A(K^0 \rightarrow \pi\pi(I)) = A_I e^{i\delta_I}, \quad A(\bar{K}^0 \rightarrow \pi\pi(I)) = A_I^* e^{i\delta_I} \quad (4)$$

where  $I$  is the isospin of  $\pi\pi$ ,  $\delta_I$  is the final-state phase shift, and  $A_I$  would be real if  $CP$  invariance held. The  $CP$ -violating observables are usually expressed in terms of  $\epsilon$  and  $\epsilon'$  defined by

$$\eta_{+-} = \epsilon + \epsilon', \quad \eta_{00} = \epsilon - 2\epsilon' \quad (5a)$$

One can then show [2]

$$\epsilon = \tilde{\epsilon} + i (\text{Im } A_0 / \text{Re } A_0), \quad (5b)$$

$$\sqrt{2}\epsilon' = ie^{i(\delta_2 - \delta_0)} (\text{Re } A_2 / \text{Re } A_0) (\text{Im } A_2 / \text{Re } A_2 - \text{Im } A_0 / \text{Re } A_0), \quad (5c)$$

$$A_L = 2\text{Re } \epsilon / (1 + |\epsilon|^2) \approx 2\text{Re } \epsilon \quad (5d)$$

In Eqs. (5a), small corrections [3] of order  $\epsilon' \times \text{Re } (A_2/A_0)$  are neglected, and Eq. (5d) assumes the  $\Delta S = \Delta Q$  rule.

The quantities  $\text{Im } A_0$ ,  $\text{Im } A_2$ , and  $\text{Im } \tilde{\epsilon}$  depend on the choice of phase convention, since one can change the phases of  $K^0$  and  $\bar{K}^0$  by a transformation of the strange quark state  $|s\rangle \rightarrow |s\rangle e^{i\alpha}$ ; of course, observables are unchanged. It is possible by a choice of phase convention to set  $\text{Im } A_0$  or  $\text{Im } A_2$  or  $\text{Im } \tilde{\epsilon}$  to zero, but none of these is zero with the usual phase conventions in the Standard Model. The choice  $\text{Im } A_0 = 0$  is called the Wu-Yang phase convention [4], in which case  $\epsilon = \tilde{\epsilon}$ . The value of  $\epsilon'$  is independent of phase convention, and a nonzero value demonstrates  $CP$  violation in the decay amplitudes, referred to

# Meson Particle Listings

$K_L^0$

as direct  $CP$  violation. The possibility that direct  $CP$  violation is essentially zero, and that  $CP$  violation occurs only in the mixing matrix, was referred to as the superweak theory [5].

By applying  $CPT$  invariance and unitarity the phase of  $\epsilon$  is given approximately by

$$\phi_\epsilon \approx \tan^{-1} \frac{2(m_{K_L} - m_{K_S})}{\Gamma_{K_S} - \Gamma_{K_L}} \approx 43.52 \pm 0.05^\circ, \quad (6a)$$

while Eq. (5c) gives the phase of  $\epsilon'$  to be

$$\phi_{\epsilon'} = \delta_2 - \delta_0 + \frac{\pi}{2} \approx 42.3 \pm 1.5^\circ, \quad (6b)$$

where the numerical value is based on an analysis of  $\pi\pi$  scattering using chiral perturbation theory [6]. The approximation in Eq. (6a) depends on the assumption that direct  $CP$  violation is very small in all  $K^0$  decays. This is expected to be good to a few tenths of a degree, as indicated by the small value of  $\epsilon'$  and of  $\eta_{+-0}$  and  $\eta_{000}$ , the  $CP$ -violation parameters in the decays  $K_S \rightarrow \pi^+\pi^-\pi^0$  [7], and  $K_S \rightarrow \pi^0\pi^0\pi^0$  [8]. The relation in Eq. (6a) is exact in the superweak theory, so this is sometimes called the superweak-phase  $\phi_{SW}$ . An important point for the analysis is that  $\cos(\phi_{\epsilon'} - \phi_\epsilon) \simeq 1$ . The consequence is that only two real quantities need be measured, the magnitude of  $\epsilon$  and the value of  $(\epsilon'/\epsilon)$ , including its sign. The measured quantity  $|\eta_{00}/\eta_{+-}|^2$  is very close to unity so that we can write

$$|\eta_{00}/\eta_{+-}|^2 \approx 1 - 6\text{Re}(\epsilon'/\epsilon) \approx 1 - 6\epsilon'/\epsilon, \quad (7a)$$

$$\text{Re}(\epsilon'/\epsilon) \approx \frac{1}{3}(1 - |\eta_{00}/\eta_{+-}|). \quad (7b)$$

From the experimental measurements in this edition of the *Review*, and the fits discussed in the next section, one finds

$$|\epsilon| = (2.228 \pm 0.011) \times 10^{-3}, \quad (8a)$$

$$\phi_\epsilon = (43.5 \pm 0.5)^\circ, \quad (8b)$$

$$\text{Re}(\epsilon'/\epsilon) \approx \epsilon'/\epsilon = (1.66 \pm 0.23) \times 10^{-3}, \quad (8c)$$

$$\phi_{+-} = (43.4 \pm 0.5)^\circ, \quad (8d)$$

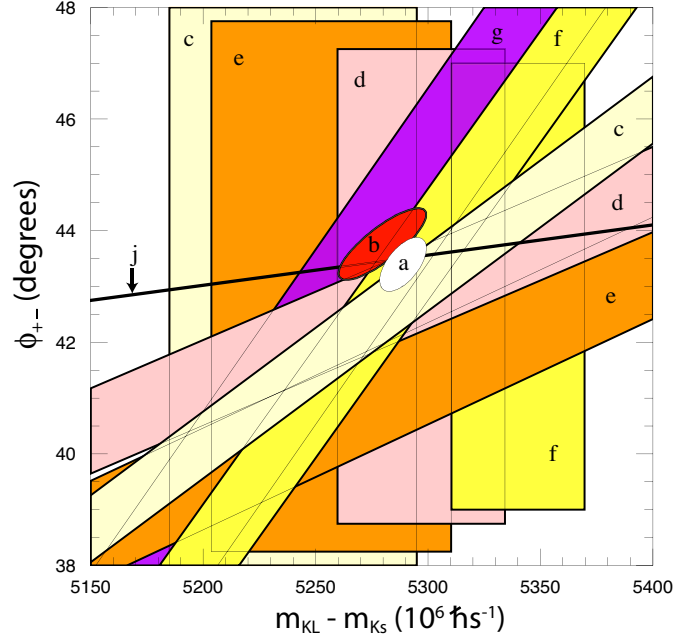
$$\phi_{00} - \phi_{+-} = (0.34 \pm 0.32)^\circ, \quad (8e)$$

$$A_L = (3.32 \pm 0.06) \times 10^{-3}. \quad (8f)$$

Direct  $CP$  violation, as indicated by  $\epsilon'/\epsilon$ , is expected in the Standard Model. However, the numerical value cannot be reliably predicted because of theoretical uncertainties [9]. The value of  $A_L$  agrees with Eq. (5d). The values of  $\phi_{+-}$  and  $\phi_{00} - \phi_{+-}$  are used to set limits on  $CPT$  violation [see “Tests of Conservation Laws”].

## 2. Fits for $K_L^0$ $CP$ -violation parameters:

In recent years,  $K_L^0$   $CP$ -violation experiments have improved our knowledge of  $CP$ -violation parameters, and their consistency with the expectations of  $CPT$  invariance and unitarity. To determine the best values of the  $CP$ -violation parameters in  $K_L^0 \rightarrow \pi^+\pi^-$  and  $\pi^0\pi^0$  decay, we make two types of fits, one for the phases  $\phi_{+-}$  and  $\phi_{00}$  jointly with  $\Delta m$  and  $\tau_S$ , and the other for the amplitudes  $|\eta_{+-}|$  and  $|\eta_{00}|$  jointly with the  $K_L^0 \rightarrow \pi\pi$  branching fractions.



**Figure 1:**  $\phi_{+-}$  vs  $\Delta m$  for experiments which do not assume  $CPT$  invariance.  $\Delta m$  measurements appear as vertical bands spanning  $\Delta m \pm 1\sigma$ , cut near the top and bottom to aid the eye. Most  $\phi_{+-}$  measurements appear as diagonal bands spanning  $\phi_{+-} \pm \sigma_\phi$ . Data are labeled by letters: “b”–FNAL KTeV, “c”–CERN CPLEAR, “d”–FNAL E773, “e”–FNAL E731, “f”–CERN, “g”–CERN NA31, and are cited in Table 1. The narrow band “j” shows  $\phi_{SW}$ . The ellipse “a” shows the  $\chi^2 = 1$  contour of the fit result.

**Table 1:** References, Document ID’s, and sources corresponding to the letter labels in the figures. The data are given in the  $\phi_{+-}$  and  $\Delta m$  sections of the  $K_L$  Listings, and the  $\tau_S$  section of the  $K_S$  Listings.

Label	Source	PDG Document ID	Ref.
a	this <i>Review</i>	OUR FIT	
b	FNAL KTeV	ABOUZAID 11	[10]
c	CERN CPLEAR	APOSTOLAKIS 99C	[11]
d	FNAL E773	SCHWINGENHEUER 95	[12]
e	FNAL E731	GIBBONS 93,93C	[13,14]
f	CERN	GEWENIGER 74B,74C	[15,16]
g	CERN NA31	CAROSI 90	[17]
h	CERN NA48	LAI 02C	[18]
i	CERN NA31	BERTANZA 97	[19]
j	this <i>Review</i>	SUPERWEAK 16	

**Fits to  $\phi_{+-}$ ,  $\phi_{00}$ ,  $\Delta\phi$ ,  $\Delta m$ , and  $\tau_S$  data:** These are joint fits to the data on  $\phi_{+-}$ ,  $\phi_{00}$ , the phase difference  $\Delta\phi = \phi_{00} - \phi_{+-}$ , the  $K_L^0 - K_S^0$  mass difference  $\Delta m$ , and the  $K_S^0$  mean life  $\tau_S$ , including the effects of correlations.

Measurements of  $\phi_{+-}$  and  $\phi_{00}$  are highly correlated with  $\Delta m$  and  $\tau_S$ . Some measurements of  $\tau_S$  are correlated with  $\Delta m$ . The correlations are given in the footnotes of the  $\phi_{+-}$  and  $\phi_{00}$  sections of the  $K_L^0$  Listings, and the  $\tau_S$  section of the  $K_S^0$  Listings.

In most cases, the correlations are quoted as 100%, *i.e.*, with the value and error of  $\phi_{+-}$  or  $\phi_{00}$  given at a fixed value of  $\Delta m$  and  $\tau_S$ , with additional terms specifying the dependence of the value on  $\Delta m$  and  $\tau_S$ . These cases lead to diagonal bands in Figs. 1 and 2. The KTeV experiment [10] quotes its results as values of  $\Delta m$ ,  $\tau_S$ ,  $\phi_\epsilon$ ,  $\text{Re}(\epsilon'/\epsilon)$ , and  $\text{Im}(\epsilon'/\epsilon)$  with correlations, leading to the ellipses labeled “b.” The correlations for the KTeV measurements are given in the  $\text{Im}(\epsilon'/\epsilon)$  section of the  $K_L^0$  Listings. For small  $|\epsilon'/\epsilon|$ ,  $\phi_{+-} \approx \phi_\epsilon + \text{Im}(\epsilon'/\epsilon)$ .

The data on  $\tau_S$ ,  $\Delta m$ , and  $\phi_{+-}$  shown in Figs. 1 and 2 are combined with data on  $\phi_{00}$  and  $\phi_{00} - \phi_{+-}$  in two fits, one without assuming *CPT*, and the other with this assumption. The results without assuming *CPT* are shown as ellipses labeled “a.” These ellipses are seen to be in good agreement with the superweak phase

$$\phi_{\text{SW}} = \tan^{-1} \left( \frac{2\Delta m}{\Delta \Gamma} \right) = \tan^{-1} \left( \frac{2\Delta m \tau_S \tau_L}{\hbar(\tau_L - \tau_S)} \right). \quad (9)$$

In Figs. 1 and 2,  $\phi_{\text{SW}}$  is shown as narrow bands labeled “j.”

Table 2 column 2, “Fit w/o *CPT*,” gives the resulting fitted parameters, while Table 3 gives the correlation matrix for this fit. The white ellipses labeled “a” in Fig. 1 and Fig. 2 are the  $\chi^2 = 1$  contours for this fit.

For experiments which have dependencies on unseen fit parameters, that is, parameters other than those shown on the x or y axis of the figure, their band positions are evaluated using the fit results and their band widths include the fitted uncertainty in the unseen parameters. This is also true for the  $\phi_{\text{SW}}$  bands.

If *CPT* invariance and unitarity are assumed, then by Eq. (6a), the phase of  $\epsilon$  is constrained to be approximately equal to

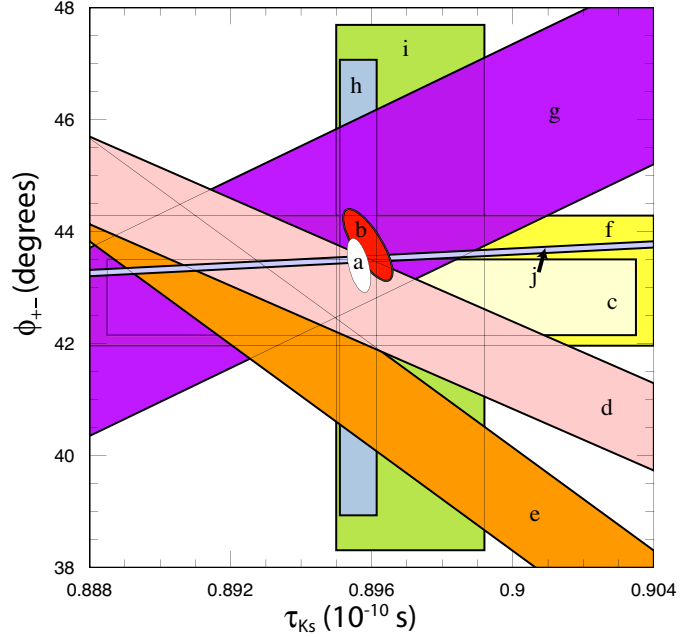
$$\begin{aligned} \phi_{\text{SW}} &= (43.50258 \pm 0.00021)^\circ + 54.1(\Delta m - 0.5289)^\circ \\ &\quad + 32.0(\tau_S - 0.89564) \end{aligned} \quad (10)$$

where we have linearized the  $\Delta m$  and  $\tau_S$  dependence of Eq. (9). The error  $\pm 0.00021$  is due to the uncertainty in  $\tau_L$ . Here  $\Delta m$  has units  $10^{10} \hbar s^{-1}$  and  $\tau_S$  has units  $10^{-10} s$ .

If in addition we use the observation that  $\text{Re}(\epsilon'/\epsilon) \ll 1$  and  $\cos(\phi_{\epsilon'} - \phi_\epsilon) \simeq 1$ , as well as the numerical value of  $\phi_{\epsilon'}$  given in Eq. (6b), then Eqs. (5a), which are sketched in Fig. 3, lead to the constraint

$$\begin{aligned} \phi_{00} - \phi_{+-} &\approx -3 \text{Im} \left( \frac{\epsilon'}{\epsilon} \right) \\ &\approx -3 \text{Re} \left( \frac{\epsilon'}{\epsilon} \right) \tan(\phi_{\epsilon'} - \phi_\epsilon) \\ &\approx 0.006^\circ \pm 0.008^\circ, \end{aligned} \quad (11)$$

so that  $\phi_{+-} \approx \phi_{00} \approx \phi_\epsilon \approx \phi_{\text{SW}}$ .

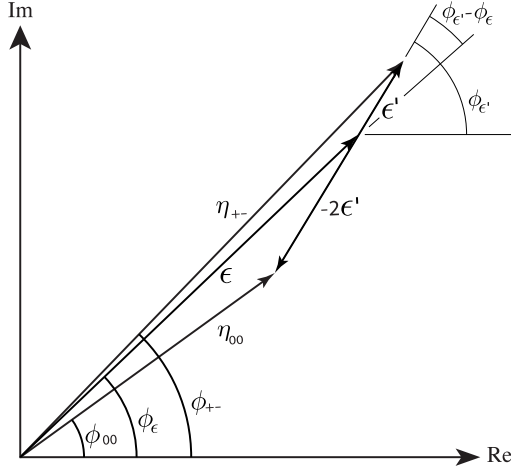


**Figure 2:**  $\phi_{+-}$  vs  $\tau_S$ .  $\tau_S$  measurements appear as vertical bands spanning  $\tau_S \pm 1\sigma$ , some of which are cut near the top and bottom to aid the eye. Most  $\phi_{+-}$  measurements appear as diagonal or horizontal bands spanning  $\phi_{+-} \pm \sigma_\phi$ . Data are labeled by letters: “b”–FNAL KTeV, “c”–CERN CPLEAR, “d”–FNAL E773, “e”–FNAL E731, “f”–CERN, “g”–CERN NA31, “h”–CERN NA48, “i”–CERN NA31, and are cited in Table 1. The narrow band “j” shows  $\phi_{\text{SW}}$ . The ellipse “a” shows the fit result’s  $\chi^2 = 1$  contour.

**Table 2:** Fit results for  $\phi_{+-}$ ,  $\Delta m$ ,  $\tau_S$ ,  $\phi_{00}$ ,  $\Delta\phi = \phi_{00} - \phi_{+-}$ , and  $\phi_\epsilon$  without and with the *CPT* assumption.

Quantity(units)	Fit w/o <i>CPT</i>	Fit w/ <i>CPT</i>
$\phi_{+-} (^\circ)$	$43.4 \pm 0.5$ (S=1.2)	$43.51 \pm 0.05$ (S=1.2)
$\Delta m (10^{10} \hbar s^{-1})$	$0.5289 \pm 0.0010$	$0.5293 \pm 0.0009$ (S=1.3)
$\tau_S (10^{-10} s)$	$0.89564 \pm 0.00033$	$0.8954 \pm 0.0004$ (S=1.1)
$\phi_{00} (^\circ)$	$43.7 \pm 0.6$ (S=1.2)	$43.52 \pm 0.05$ (S=1.3)
$\Delta\phi (^\circ)$	$0.34 \pm 0.32$	$0.006 \pm 0.014$ (S=1.7)
$\phi_\epsilon (^\circ)$	$43.5 \pm 0.5$ (S=1.3)	$43.52 \pm 0.05$ (S=1.2)
$\chi^2$	16.4	20.0
# Deg. Free.	14	16

In the fit assuming *CPT*, we constrain  $\phi_\epsilon = \phi_{\text{SW}}$  using the linear expression in Eq. (10), and constrain  $\phi_{00} - \phi_{+-}$  using Eq. (11). These constraints are inserted into the Listings with the Document ID of SUPERWEAK 16. Some additional data for which the authors assumed *CPT* are added to this fit or substitute for other less precise data for which the authors did not make this assumption. See the Listings for details.



**Figure 3:** Sketch of Eqs. (5a). Not to scale.

The results of this fit are shown in Table 2, column 3, “Fit w/*CPT*,” and the correlation matrix is shown in Table 4. The  $\Delta m$  precision is improved by the *CPT* assumption.

**Table 3:** Correlation matrix for the results of the fit without the *CPT* assumption

	$\phi_{+-}$	$\Delta m$	$\tau_S$	$\phi_{00}$	$\Delta\phi$	$\phi_\epsilon$
$\phi_{+-}$	1.000	0.596	-0.488	0.827	-0.040	0.976
$\Delta m$	0.596	1.000	-0.572	0.487	-0.035	0.580
$\tau_S$	-0.488	-0.572	1.000	-0.423	-0.014	-0.484
$\phi_{00}$	0.827	0.487	-0.423	1.000	0.529	0.929
$\Delta\phi$	-0.040	-0.035	-0.014	0.529	1.000	0.178
$\phi_\epsilon$	0.976	0.580	-0.484	0.929	0.178	1.000

**Table 4:** Correlation matrix for the results of the fit with the *CPT* assumption

	$\phi_{+-}$	$\Delta m$	$\tau_S$	$\phi_{00}$	$\Delta\phi$	$\phi_\epsilon$
$\phi_{+-}$	1.000	0.972	-0.311	0.957	-0.105	0.995
$\Delta m$	0.972	1.000	-0.509	0.958	-0.007	0.977
$\tau_S$	-0.311	-0.509	1.000	-0.306	0.004	-0.312
$\phi_{00}$	0.957	0.958	-0.306	1.000	0.189	0.981
$\Delta\phi$	-0.105	-0.007	0.004	0.189	1.000	-0.006
$\phi_\epsilon$	0.995	0.977	-0.312	0.981	-0.006	1.000

#### Fits for $\epsilon'/\epsilon$ , $|\eta_{+-}|$ , $|\eta_{00}|$ , and $B(K_L \rightarrow \pi\pi)$

We list measurements of  $|\eta_{+-}|$ ,  $|\eta_{00}|$ ,  $|\eta_{00}/\eta_{+-}|$ , and  $\epsilon'/\epsilon$ . Independent information on  $|\eta_{+-}|$  and  $|\eta_{00}|$  can be obtained from measurements of the  $K_L^0$  and  $K_S^0$  lifetimes ( $\tau_L$ ,  $\tau_S$ ), and branching ratios ( $B$ ) to  $\pi\pi$ , using the relations

$$|\eta_{+-}| = \left[ \frac{B(K_L^0 \rightarrow \pi^+\pi^-)}{\tau_L} \frac{\tau_S}{B(K_S^0 \rightarrow \pi^+\pi^-)} \right]^{1/2}, \quad (12a)$$

$$|\eta_{00}| = \left[ \frac{B(K_L^0 \rightarrow \pi^0\pi^0)}{\tau_L} \frac{\tau_S}{B(K_S^0 \rightarrow \pi^0\pi^0)} \right]^{1/2}. \quad (12b)$$

For historical reasons, the branching ratio fits and the *CP*-violation fits are done separately, but we want to include the influence of  $|\eta_{+-}|$ ,  $|\eta_{00}|$ ,  $|\eta_{00}/\eta_{+-}|$ , and  $\epsilon'/\epsilon$  measurements on  $B(K_L^0 \rightarrow \pi^+\pi^-)$  and  $B(K_L^0 \rightarrow \pi^0\pi^0)$  and vice versa. We approximate a global fit to all of these measurements by first performing two independent fits: 1) BRFIT, a fit to the  $K_L^0$  branching ratios, rates, and mean life, and 2) ETAFIT, a fit to the  $|\eta_{+-}|$ ,  $|\eta_{00}|$ ,  $|\eta_{+-}/\eta_{00}|$ , and  $\epsilon'/\epsilon$  measurements. The results from fit 1, along with the  $K_S^0$  values from this edition, are used to compute values of  $|\eta_{+-}|$  and  $|\eta_{00}|$ , which are included as measurements in the  $|\eta_{00}|$  and  $|\eta_{+-}|$  sections with a document ID of BRFIT 16. Thus, the fit values of  $|\eta_{+-}|$  and  $|\eta_{00}|$  given in this edition include both the direct measurements and the results from the branching ratio fit.

The process is reversed in order to include the direct  $|\eta|$  measurements in the branching ratio fit. The results from fit 2 above (before including BRFIT 16 values) are used along with the  $K_L^0$  and  $K_S^0$  mean lives and the  $K_S^0 \rightarrow \pi\pi$  branching fractions to compute the  $K_L^0$  branching ratio  $\Gamma(K_L^0 \rightarrow \pi^0\pi^0)/\Gamma(K_L^0 \rightarrow \pi^+\pi^-)$ . This branching ratio value is included as a measurement in the branching ratio section with a document ID of ETAFIT 16. Thus, the  $K_L^0$  branching ratio fit values in this edition include the results of the direct measurement of  $|\eta_{00}/\eta_{+-}|$  and  $\epsilon'/\epsilon$ . Most individual measurements of  $|\eta_{+-}|$  and  $|\eta_{00}|$  enter our fits directly via the corresponding measurements of  $\Gamma(K_L^0 \rightarrow \pi^+\pi^-)/\Gamma(\text{total})$  and  $\Gamma(K_L^0 \rightarrow \pi^0\pi^0)/\Gamma(\text{total})$ , and those that do not have too large errors to have any influence on the fitted values of these branching ratios. A more detailed discussion of these fits is given in the 1990 edition of this *Review* [20].

#### References

1. K. Kleinknecht, “Uncovering *CP* violation: experimental clarification in the neutral *K* meson and *B* meson systems,” *Springer Tracts in Modern Physics*, vol. 195 (Springer Verlag 2003).
2. B. Winstein and L. Wolfenstein, *Rev. Mod. Phys.* **65**, 1113 (1993).
3. M.S. Sozzi, *Eur. Phys. J.* **C36**, 37 (2004).
4. T.T. Wu and C.N. Yang, *Phys. Rev. Lett.* **13**, 380 (1964).
5. L. Wolfenstein, *Phys. Rev. Lett.* **13**, 562 (1964);  
L. Wolfenstein, *Comm. Nucl. Part. Phys.* **21**, 275 (1994).
6. G. Colangelo, J. Gasser, and H. Leutwyler, *Nucl. Phys.* **B603**, 125 (2001).
7. R. Adler *et al.*, (CPLEAR Collab.), *Phys. Lett.* **B407**, 193 (1997);  
P. Bloch, *Proceedings of Workshop on K Physics* (Orsay 1996), ed. L. Ikonomidou-Fayard, Edition Frontieres, Gif-sur-Yvette, France (1997) p. 307.
8. A. Lai *et al.*, *Phys. Lett.* **B610**, 165 (2005).
9. G. Buchalla, A.J. Buras, and M.E. Lautenbacher, *Rev. Mod. Phys.* **68**, 1125 (1996);  
S. Bosch *et al.*, *Nucl. Phys.* **B565**, 3 (2000);

S. Bertolini, M. Fabrichesi, and J.O. Egg, Rev. Mod. Phys. **72**, 65 (2000).

10. E. Abouzaid *et al.*, Phys. Rev. **D83**, 092001 (2011).
11. A. Apostolakis *et al.*, Phys. Lett. **B458**, 545 (1999).
12. B. Schwingenheuer *et al.*, Phys. Rev. Lett. **74**, 4376 (1995).
13. L.K. Gibbons *et al.*, Phys. Rev. Lett. **70**, 1199 (1993) and footnote in Ref. 12.
14. L.K. Gibbons, Thesis, RX-1487, Univ. of Chicago, 1993.
15. C. Geweniger *et al.*, Phys. Lett. **48B**, 487 (1974).
16. C. Geweniger *et al.*, Phys. Lett. **52B**, 108 (1974).
17. R. Carosi *et al.*, Phys. Lett. **B237**, 303 (1990).
18. A. Lai *et al.*, Phys. Lett. **B537**, 28 (2002).
19. L. Bertanza *et al.*, Z. Phys. **C73**, 629 (1997).
20. J.J. Hernandez *et al.*, Particle Data Group, Phys. Lett. **B239**, 1 (1990).

### CP-VIOLATION PARAMETERS IN $K_L^0$ DECAYS

#### CHARGE ASYMMETRY IN $K_S^0$ DECAYS

Such asymmetry violates *CP*. It is related to  $\text{Re}(\epsilon)$ .

#### $A_L$ = weighted average of $A_L(\mu)$ and $A_L(e)$

In previous editions and in the literature the symbol used for this asymmetry was  $\delta_L^{\pm}$  or  $\delta$ . We use  $A_L$  for consistency with  $B^0$  asymmetry notation and with recent  $K_S^0$  notation.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.332±0.006 OUR AVERAGE</b>		Includes data from the 2 datablocks that follow this one.		
0.333±0.050	33M	WILLIAMS	73	ASPK $K_{\mu 3} + K_{e 3}$

#### $A_L(\mu) = [\Gamma(\pi^- \mu^+ \nu_\mu) - \Gamma(\pi^+ \mu^- \bar{\nu}_\mu)]/\text{SUM}$

Only the combined value below is put into the Meson Summary Table.

VALUE (%)	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

#### **0.304±0.025 OUR AVERAGE**

0.313±0.029	15M	GEWENIGER	74	ASPK
0.278±0.051	7.7M	PICCONI	72	ASPK
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.60 ±0.14	4.1M	MCCARTHY	73	CNTR
0.57 ±0.17	1M	<sup>1</sup> PACIOTTI	69	OSPK
0.403±0.134	1M	<sup>1</sup> DORFAN	67	OSPK

<sup>1</sup> PACIOTTI 69 is a reanalysis of DORFAN 67 and is corrected for  $\mu^+ \mu^-$  range difference in MCCARTHY 72.

#### $A_L(e) = [\Gamma(\pi^- e^+ \nu_e) - \Gamma(\pi^+ e^- \bar{\nu}_e)]/\text{SUM}$

Only the combined value below is put into the Meson Summary Table.

VALUE (%)	EVTS	DOCUMENT ID	TECN
The data in this block is included in the average printed for a previous datablock.			

#### **0.334 ±0.007 OUR AVERAGE**

0.3322±0.0058±0.0047	298M	ALAVI-HARATI	02	
0.341 ±0.018	34M	GEWENIGER	74	ASPK
0.318 ±0.038	40M	FITCH	73	ASPK
0.346 ±0.033	10M	MARX	70	CNTR
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.36 ±0.18	600k	ASHFORD	72	ASPK
0.246 ±0.059	10M	<sup>1</sup> SAAL	69	CNTR
0.224 ±0.036	10M	<sup>1</sup> BENNETT	67	CNTR

<sup>1</sup> SAAL 69 is a reanalysis of BENNETT 67.

### PARAMETERS FOR $K_L^0 \rightarrow 2\pi$ DECAY

$$\eta_{+-} = A(K_L^0 \rightarrow \pi^+ \pi^-) / A(K_S^0 \rightarrow \pi^+ \pi^-)$$

$$\eta_{00} = A(K_L^0 \rightarrow \pi^0 \pi^0) / A(K_S^0 \rightarrow \pi^0 \pi^0)$$

The fitted values of  $|\eta_{+-}|$  and  $|\eta_{00}|$  given below are the results of a fit to  $|\eta_{+-}|$ ,  $|\eta_{00}|$ ,  $|\eta_{00}/\eta_{+-}|$ , and  $\text{Re}(\epsilon'/\epsilon)$ . Independent information on  $|\eta_{+-}|$  and  $|\eta_{00}|$  can be obtained from the fitted values of the  $K_L^0 \rightarrow \pi\pi$  and  $K_S^0 \rightarrow \pi\pi$  branching ratios and the  $K_L^0$  and  $K_S^0$  lifetimes. This information is included as data in the  $|\eta_{+-}|$  and  $|\eta_{00}|$  sections with a Document ID "BRFIT." See the note "CP violation in  $K_L^0$  decays" above for details.

$$|\eta_{00}| = |A(K_L^0 \rightarrow 2\pi^0) / A(K_S^0 \rightarrow 2\pi^0)|$$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.220±0.011 OUR FIT</b>	Error includes scale factor of 1.8.		
<b>2.243±0.014</b>	BRFIT	16	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.47 ±0.31 ±0.24	ANGELOPO...	98	CPLR
2.49 ±0.40	<sup>1</sup> ADLER	96B	CPLR Sup. by ANGELOPOULOS 98
2.33 ±0.18	CHRISTENS...	79	ASPK
2.71 ±0.37	<sup>2</sup> WOLFF	71	OSPK Cu reg., $4\gamma$ 's
2.95 ±0.63	<sup>2</sup> CHOLLET	70	OSPK Cu reg., $4\gamma$ 's

<sup>1</sup> Error is statistical only.

<sup>2</sup> CHOLLET 70 gives  $|\eta_{00}| = (1.23 \pm 0.24) \times (\text{regeneration amplitude, 2 GeV/c Cu})/10000\text{mb}$ . WOLFF 71 gives  $|\eta_{00}| = (1.13 \pm 0.12) \times (\text{regeneration amplitude, 2 GeV/c Cu})/10000\text{mb}$ . We compute both  $|\eta_{00}|$  values for (regeneration amplitude, 2 GeV/c Cu) =  $24 \pm 2\text{mb}$ . This regeneration amplitude results from averaging over FAISSNER 69, extrapolated using optical-model calculations of Bohm *et al.*, Physics Letters **27B** 594 (1968) and the data of BALATS 71. (From H. Faissner, private communication).

$$|\eta_{+-}| = |A(K_L^0 \rightarrow \pi^+ \pi^-) / A(K_S^0 \rightarrow \pi^+ \pi^-)|$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.232±0.011 OUR FIT</b>	Error includes scale factor of 1.8.			
<b>2.226±0.007</b>		BRFIT	16	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.223±0.012		<sup>1</sup> LAI	07	NA48
2.219±0.013		<sup>2</sup> AMBROSINO	06F	KLOE
2.228±0.010		<sup>3</sup> ALEXOPOU...	04	KTEV
2.286±0.023±0.026	70M	<sup>4</sup> APOSTOLA...	99C	CPLR $K^0\text{-}\bar{K}^0$ asymmetry
2.310±0.043±0.031		<sup>5</sup> ADLER	95B	CPLR $K^0\text{-}\bar{K}^0$ asymmetry
2.32 ±0.14 ±0.03	10 <sup>5</sup>	ADLER	92B	CPLR $K^0\text{-}\bar{K}^0$ asymmetry
2.30 ±0.035		GEWENIGER	74B	ASPK

<sup>1</sup> Value obtained from the NA48 measurements of  $\Gamma(K_L^0 \rightarrow \pi^+ \pi^-)/\Gamma(K_L^0 \rightarrow \pi e \nu_e)$  and  $\tau_{K_S^0}$  and KLOE measurements of  $B(K_S^0 \rightarrow \pi^+ \pi^-)$  and  $\tau_{K_L^0}$ .  $\Gamma(K_L^0 \rightarrow \pi^+ \pi^-)$  is defined to include the inner bremsstrahlung component  $\Gamma(K_L^0 \rightarrow \pi^+ \pi^- \gamma(\text{IB}))$  but

exclude the direct emission component  $B(K_S^0 \rightarrow \pi^+ \pi^- (\text{DE}))$ . Their  $|\eta_{+-}|$  value is not directly used in our fit, but enters the fit via their branching ratio and lifetime measurements.

<sup>2</sup> AMBROSINO 06F uses KLOE branching ratios and  $\tau_L$  together with  $\tau_S$  from PDG 04. Their  $|\eta_{+-}|$  value is not directly used in our fit, but enters the fit via their branching ratio and lifetime measurements.

<sup>3</sup> ALEXOPOULOS 04  $|\eta_{+-}|$  uses their  $K_L^0 \rightarrow \pi\pi$  branching fractions,  $\tau_S = (0.8963 \pm 0.0005) \times 10^{-10}\text{s}$  from the average of KTeV and NA48  $\tau_S$  measurements, and assumes that  $\Gamma(K_L^0 \rightarrow \pi e \nu_e) = \Gamma(K_L^0 \rightarrow \pi \ell \nu_\ell)$  giving  $B(K_S^0 \rightarrow \pi \ell \nu_\ell) = 0.118\%$ . Their  $\eta_{+-}$  is not directly used in our fit, but enters our fit via their branching ratio measurements.

<sup>4</sup> APOSTOLAKIS 99C report  $(2.264 \pm 0.023 \pm 0.026 + 9.1[\tau_S - 0.8934]) \times 10^{-3}$ . We evaluate for our 2006 best value  $\tau_S = (0.8958 \pm 0.0005) \times 10^{-10}\text{s}$ .

<sup>5</sup> ADLER 95B report  $(2.312 \pm 0.043 \pm 0.030 - 1[\Delta m - 0.5274] + 9.1[\tau_S - 0.8926]) \times 10^{-3}$ . We evaluate for our 1996 best values  $\Delta m = (0.5304 \pm 0.0014) \times 10^{-10}\text{fs}^{-1}$  and  $\tau_S = (0.8927 \pm 0.0009) \times 10^{-10}\text{s}$ . Superseded by APOSTOLAKIS 99C.

$$|\epsilon| = (2|\eta_{+-}| + |\eta_{00}|)/3$$

This expression is a very good approximation, good to about one part in  $10^{-4}$  because of the small measured value of  $\phi_{00} - \phi_{+-}$  and small theoretical ambiguities.

VALUE (units $10^{-3}$ )	DOCUMENT ID
<b>2.228±0.011 OUR FIT</b>	Error includes scale factor of 1.8.

$$|\eta_{00}/\eta_{+-}|$$

VALUE	EVTS	DOCUMENT ID	TECN
<b>0.9950±0.0007 OUR FIT</b>	Error includes scale factor of 1.6.		
<b>0.9930±0.0020 OUR AVERAGE</b>			
0.9931±0.0020		<sup>1,2</sup> BARR	93D NA 31
0.9904±0.0084±0.0036		<sup>3</sup> WOODS	88 E731
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.9939±0.0013±0.0015	1M	<sup>1</sup> BARR	93D NA 31
0.9899±0.0020±0.0025		<sup>1</sup> BURKHARDT	88 NA 31

<sup>1</sup> This is the square root of the ratio  $R$  given by BURKHARDT 88 and BARR 93D.

<sup>2</sup> This is the combined results from BARR 93D and BURKHARDT 88, taking into account a common systematic uncertainty of 0.0014.

<sup>3</sup> We calculate  $|\eta_{00}/\eta_{+-}| = 1 - 3(\epsilon'/\epsilon)$  from WOODS 88 ( $\epsilon'/\epsilon$ ) value.

$$\text{Re}(\epsilon'/\epsilon) = (1 - |\eta_{00}/\eta_{+-}|)/3$$

We have neglected terms of order  $\omega \cdot \text{Re}(\epsilon'/\epsilon)$ , where  $\omega = \text{Re}(A_2)/\text{Re}(A_0) \simeq 1/22$ . If included, this correction would lower  $\text{Re}(\epsilon'/\epsilon)$  by about  $0.04 \times 10^{-3}$ . See SOZZI 04.

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.66 ±0.23 OUR FIT</b>	Error includes scale factor of 1.6.		
<b>1.68 ±0.20 OUR AVERAGE</b>	Error includes scale factor of 1.4. See the ideogram below.		
1.92 ±0.21	<sup>1</sup> ABOUZAIID	11	KTEV Assuming CPT
1.47 ±0.22	BATLEY	02	NA48
0.74 ±0.52 ±0.29	GIBBONS	93B	E731
• • • We use the following data for averages but not for fits. • • •			
2.3 ±0.65	<sup>2,3</sup> BARR	93D	NA 31

# Meson Particle Listings

 $K_L^0$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.110 ± 0.343	1.4	ABOUZAID	11	KTEV	Not assuming <i>CPT</i>
2.07 ± 0.28		ALAVI-HARATI	03	KTEV	In ABOUZAID 11
1.53 ± 0.26		LAI	01C	NA48	Incl. in BATLEY 02
2.80 ± 0.30 ± 0.28		ALAVI-HARATI	99D	KTEV	In ALAVI-HARATI 03
1.85 ± 0.45 ± 0.58		FANTI	99C	NA48	In LAI 01C
2.0 ± 0.7	5	BARR	93D	NA31	
-0.4 ± 1.4 ± 0.6		PATTERSON	90	E731	In GIBBONS 93B
3.3 ± 1.1	5	BURKHARDT	88	NA31	
3.2 ± 2.8 ± 1.2	2	WOODS	88	E731	

<sup>1</sup> The two ABOUZAID 11 values use the same data. The fits are performed with and without *CPT* invariance requirement.

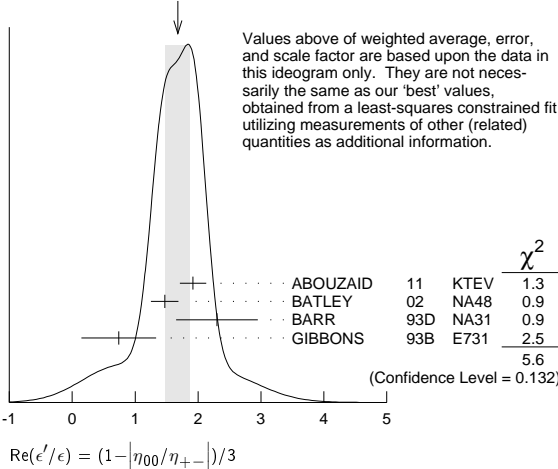
<sup>2</sup> These values are derived from  $|\eta_{00}/\eta_{+-}|$  measurements. They enter the average in this section but enter the fit via the  $|\eta_{00}/\eta_{+-}|$  only.

<sup>3</sup> This is the combined results from BARR 93D and BURKHARDT 88, taking into account their common systematic uncertainty.

<sup>4</sup> We use ABOUZAID 11  $\text{Re}(\epsilon'/\epsilon)$  value with *CPT* assumption in our fits for  $|\eta_{+-}|$ ,  $|\eta_{00}|$ , and  $\text{Re}(\epsilon'/\epsilon)$ .

<sup>5</sup> These values are derived from  $|\eta_{00}/\eta_{+-}|$  measurements.

WEIGHTED AVERAGE  
1.68 ± 0.20 (Error scaled by 1.4)



## $\phi_{+-}$ , PHASE of $\eta_{+-}$

The dependence of the phase on  $\Delta m$  and  $\tau_S$  is given for each experiment in the comments below, where  $\Delta m$  is the  $K_L^0 - K_S^0$  mass difference in units  $10^{10} \text{ h s}^{-1}$  and  $\tau_S$  is the  $K_S$  mean life in units  $10^{-10} \text{ s}$ . We also give the regeneration phase  $\phi_f$  in the comments below.

OUR FIT is described in the note on “*CP* violation in  $K_L$  decays” in the  $K_L^0$  Particle Listings. Most experiments in this section are included in both the “Not Assuming *CPT*” and “Assuming *CPT*” fits. In the latter fit, they have little direct influence on  $\phi_{+-}$  because their errors are large compared to that assuming *CPT*, but they influence  $\Delta m$  and  $\tau_S$  through their dependencies on these parameters, which are given in the footnotes.

VALUE (°)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>43.51 ± 0.05 OUR FIT</b>				Error includes scale factor of 1.2. Assuming <i>CPT</i>
<b>43.4 ± 0.5 OUR FIT</b>				Error includes scale factor of 1.2. Not assuming <i>CPT</i>
42.9 ± 0.6 ± 0.3	70M	<sup>1</sup> APOSTOLAKIS... 99c	CPLR	$K^0 \bar{K}^0$ asymmetry
42.9 ± 0.8 ± 0.2		<sup>2,3</sup> SCHWINGENHEUER... 95	E773	$\text{CH}_{1,1}$ regenerator
41.4 ± 0.9 ± 0.2		<sup>3,4</sup> GIBBONS 93	E731	$B_4C$ regenerator
44.5 ± 1.6 ± 0.6		<sup>5</sup> CAROSI 90	NA31	Vacuum regen.
43.3 ± 1.0 ± 0.5		<sup>6</sup> GEWENIGER 74B	ASPK	Vacuum regen.
• • • We do not use the following data for averages, fits, limits, etc. • • •				
43.76 ± 0.64		<sup>7</sup> ABOUZAID 11	KTEV	Not assuming <i>CPT</i>
44.12 ± 0.72 ± 1.20		<sup>8</sup> ALAVI-HARATI 03	KTEV	Not assuming <i>CPT</i>
42.5 ± 0.4 ± 0.3		<sup>9,10</sup> ADLER 96C	RVUE	
43.4 ± 1.1 ± 0.3		<sup>11</sup> ADLER 95B	CPLR	$K^0 \bar{K}^0$ asymmetry
42.3 ± 4.4 ± 1.4	100k	<sup>12</sup> ADLER 92B	CPLR	$K^0 \bar{K}^0$ asymmetry
47.7 ± 2.0 ± 0.9		<sup>3,13</sup> KARLSSON 90	E731	
44.3 ± 2.8 ± 0.2		<sup>14</sup> CARITHERS 75	SPEC	C regenerator

<sup>1</sup> APOSTOLAKIS 99c measures  $\phi_{+-} = (43.19 \pm 0.53 \pm 0.28) + 300 [\Delta m - 0.5301] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> SCHWINGENHEUER 95 measures  $\phi_{+-} = (43.53 \pm 0.76) + 173 [\Delta m - 0.5282] - 275 [\tau_S - 0.8926] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>3</sup> These experiments measure  $\phi_{+-} - \phi_f$  and calculate the regeneration phase from the power law momentum dependence of the regeneration amplitude using analyticity and dispersion relations. SCHWINGENHEUER 95 [GIBBONS 93] includes a systematic error of  $0.35^\circ$  [0.5°] for uncertainties in their modeling of the regeneration amplitude.

<sup>4</sup> GIBBONS 93 measures  $\phi_{+-} = (42.21 \pm 0.9) + 189 [\Delta m - 0.5257] - 460 [\tau_S - 0.8922] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values. This is actually reported in SCHWINGENHEUER 95, footnote 8. GIBBONS 93 reports  $\phi_{+-} = (42.2 \pm 1.4)^\circ$ . They measure  $\phi_{+-} - \phi_f$  and calculate the regeneration phase  $\phi_f$  from the power law momentum dependence of the regeneration amplitude using analyticity. An error of  $0.6^\circ$  is included for possible uncertainties in the regeneration phase.

<sup>5</sup> CAROSI 90 measures  $\phi_{+-} = (46.9 \pm 1.4 \pm 0.7) + 579 [\Delta m - 0.5351] + 303 [\tau_S - 0.8922] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>6</sup> GEWENIGER 74B measures  $\phi_{+-} = (49.4 \pm 1.0) + 565 [\Delta m - 0.540] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>7</sup> Not independent of other phase parameters reported in ABOUZAID 11.

<sup>8</sup> ALAVI-HARATI 03  $\phi_{+-}$  is correlated with their  $\Delta m = m_{K_L^0} - m_{K_S^0}$  and  $\tau_{K_S}$  measurements in the  $K_L^0$  and  $K_S^0$  sections respectively. The correlation coefficients are  $\rho(\phi_{+-}, \Delta m) = +0.955$ ,  $\rho(\phi_{+-}, \tau_S) = -0.871$ , and  $\rho(\tau_S, \Delta m) = -0.840$ . *CPT* is not assumed. Uses scintillator Pb regenerator. Superseded by ABOUZAID 11.

<sup>9</sup> ADLER 96C measures  $\phi_{+-} = (43.82 \pm 0.41) + 339 [\Delta m - 0.5307] - 252 [\tau_S - 0.8922] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>10</sup> ADLER 96C is the result of a fit which includes nearly the same data as entered into the “OUR FIT” value in the 1996 edition of this Review (Physical Review **D54** 1 (1996)).

<sup>11</sup> ADLER 95B measures  $\phi_{+-} = (42.7 \pm 0.9 \pm 0.6) + 316 [\Delta m - 0.5274] + 30 [\tau_S - 0.8926] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>12</sup> ADLER 92B quote separately two systematic errors:  $\pm 0.4$  from their experiment and  $\pm 1.0$  degrees due to the uncertainty in the value of  $\Delta m$ .

<sup>13</sup> KARLSSON 90 systematic error does not include regeneration phase uncertainty.

<sup>14</sup> CARITHERS 75 measures  $\phi_{+-} = (45.5 \pm 2.8) + 224 [\Delta m - 0.5348] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.  $\phi_f = -40.9 \pm 2.6^\circ$ .

## $\phi_{00}$ , PHASE of $\eta_{00}$

See comment in  $\phi_{+-}$  header above for treatment of  $\Delta m$  and  $\tau_S$  dependence, as well as for the inclusion of data in both the “Assuming *CPT*” and “Not Assuming *CPT*” fits.

OUR FIT is described in the note on “*CP* violation in  $K_L$  decays” in the  $K_L^0$  Particle Listings.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>43.52 ± 0.05 OUR FIT</b>			Error includes scale factor of 1.3. Assuming <i>CPT</i>
<b>43.7 ± 0.6 OUR FIT</b>			Error includes scale factor of 1.2. Not assuming <i>CPT</i>
44.5 ± 2.3 ± 0.5	<sup>1</sup> CAROSI 90	NA31	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
44.06 ± 0.68	<sup>2</sup> ABOUZAID 11	KTEV	Not assuming <i>CPT</i>
41.7 ± 5.9 ± 0.2	<sup>3</sup> ANGELOPOULOS... 98	CPLR	
50.8 ± 7.1 ± 1.7	<sup>4</sup> ADLER 96B	CPLR	Sup. by ANGELOPOULOS 98
47.4 ± 1.4 ± 0.9	<sup>5</sup> KARLSSON 90	E731	

<sup>1</sup> CAROSI 90 measures  $\phi_{00} = (47.1 \pm 2.1 \pm 1.0) + 579 [\Delta m - 0.5351] + 252 [\tau_S - 0.8922] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> Not independent of other phase parameters reported in ABOUZAID 11.

<sup>3</sup> ANGELOPOULOS 98 measures  $\phi_{00} = (42.0 \pm 5.6 \pm 1.9) + 240 [\Delta m - 0.5307] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values. The  $\tau_S$  dependence is negligible.

<sup>4</sup> ADLER 96B identified initial neutral kaon individually as being a  $K^0$  or a  $\bar{K}^0$ . The systematic uncertainty is  $\pm 1.5^\circ$  combined in quadrature with  $\pm 0.8^\circ$  due to  $\Delta m$ .

<sup>5</sup> KARLSSON 90 systematic error does not include regeneration phase uncertainty.

## $\phi_e = (2\phi_{+-} + \phi_{00})/3$

This expression is a very good approximation, good to about  $10^{-3}$  degrees because of the small measured values of  $\phi_{00} - \phi_{+-}$  and  $\text{Re } \epsilon'/\epsilon$ , and small theoretical ambiguities.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>43.52 ± 0.05 OUR FIT</b>			Error includes scale factor of 1.2. Assuming <i>CPT</i>
<b>43.5 ± 0.5 OUR FIT</b>			Error includes scale factor of 1.3. Not assuming <i>CPT</i>
43.5164 ± 0.0002 ± 0.0518	<sup>1</sup> SUPERWEAK 16		Assuming <i>CPT</i>
43.86 ± 0.63	<sup>2</sup> ABOUZAID 11	KTEV	Not assuming <i>CPT</i>

<sup>1</sup> SUPERWEAK 16 is a fake measurement used to impose the *CPT* or Superweak constraint  $\phi_{+-} = \phi_{SW} = \tan^{-1}[2 \frac{\Delta m}{h} (\frac{\tau_S \tau_L}{\tau_L - \tau_S})]$ . This “measurement” is linearized using values near the PDG 04 edition values of  $\Delta m$ ,  $\tau_S$  and  $\tau_L$ , and then adjusted to our current values as described in the following “measurement”. SUPERWEAK 16 measures  $\phi_e = (43.50258 \pm 0.00021) + 54.1 [\Delta m - 0.5289] + 32.0 [\tau_S - 0.89564] (^\circ)$ . We have adjusted the measurement to use our best values of ( $\Delta m = 0.5293 \pm 0.0009$ ) ( $10^{10} \text{ h s}^{-1}$ ), ( $\tau_S = 0.8954 \pm 0.0004$ ) ( $10^{-10} \text{ s}$ ). Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> ABOUZAID 11 uses the full KTeV dataset collected in 1996, 1997, and 1999. See  $\text{Im}(\epsilon'/\epsilon)$  section for correlation information.



See key on page 601

# Meson Particle Listings

$K_L^0$

$$\text{Im}(\epsilon'/\epsilon) = -(\phi_{00} - \phi_{+-})/3$$

For small  $|\epsilon'/\epsilon|$ ,  $\text{Im}(\epsilon'/\epsilon)$  is related to the phases of  $\eta_{00}$  and  $\eta_{+-}$  by the above expression.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-0.002 ± 0.005 OUR FIT</b>			Error includes scale factor of 1.7. Assuming <i>CPT</i>
<b>-0.11 ± 0.11 OUR FIT</b>			Not assuming <i>CPT</i>
<b>-0.0985 ± 0.1157</b>	1 ABOUZAIID	11	KTEV Not assuming <i>CPT</i>

<sup>1</sup> ABOUZAIID 11 uses the full KTeV dataset collected in 1996, 1997, and 1999. The fit has  $\Delta m$ ,  $\tau_S$ ,  $\phi_\epsilon$ ,  $\text{Re}(\epsilon'/\epsilon)$ , and  $\text{Im}(\epsilon'/\epsilon)$  as free parameters. The reported value of  $\text{Im}(\epsilon'/\epsilon) = (-17.20 \pm 20.20) \times 10^{-4}$  rad. The correlation coefficients are  $\rho(\phi_\epsilon, \Delta m) = 0.828$ ,  $\rho(\phi_\epsilon, \tau_S) = -0.765$ ,  $\rho(\Delta m, \tau_S) = -0.858$ ,  $\rho(\text{Im}(\epsilon'/\epsilon), \phi_\epsilon) = -0.041$ ,  $\rho(\text{Im}(\epsilon'/\epsilon), \Delta m) = 0.026$ ,  $\rho(\text{Im}(\epsilon'/\epsilon), \tau_S) = -0.010$ .

## DECAY-PLANE ASYMMETRY IN $\pi^+\pi^-e^+e^-$ DECAYS

This is the *CP*-violating asymmetry

$$A = \frac{N_{\sin\phi\cos\phi>0.0} - N_{\sin\phi\cos\phi<0.0}}{N_{\sin\phi\cos\phi>0.0} + N_{\sin\phi\cos\phi<0.0}}$$

where  $\phi$  is the angle between the  $e^+e^-$  and  $\pi^+\pi^-$  planes in the  $K_L^0$  rest frame.

## CP ASYMMETRY $A$ in $K_L^0 \rightarrow \pi^+\pi^-e^+e^-$

VALUE (%)	DOCUMENT ID	TECN
<b>13.7 ± 1.5 OUR AVERAGE</b>		
13.6 ± 1.4 ± 1.5	ABOUZAIID	06 KTEV
14.2 ± 3.0 ± 1.9	LAI	03c NA48
13.6 ± 2.5 ± 1.2	ALAVI-HARATI00B	KTEV

## PARAMETERS FOR $e^+e^-e^+e^-$ DECAYS

These are the *CP*-violating parameters in the  $\phi$  distribution, where  $\phi$  is the angle between the planes of the two  $e^+e^-$  pairs in the kaon rest frame:

$$d\Gamma/d\phi \propto 1 + \beta_{CP} \cos(2\phi) + \gamma_{CP} \sin(2\phi)$$

## $\beta_{CP}$ from $K_L^0 \rightarrow e^+e^-e^+e^-$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.19 ± 0.07 OUR AVERAGE</b>				
-0.13 ± 0.10 ± 0.03	200	<sup>1</sup> LAI	05B NA48	
-0.23 ± 0.09 ± 0.02	441	ALAVI-HARATI01D	KTEV	$M_{ee} > 8 \text{ MeV}/c^2$

<sup>1</sup> LAI 05B obtains  $\beta_{CP} = -0.13 \pm 0.10$  (stat) if  $\gamma_{CP} = 0$  is assumed.

## $\gamma_{CP}$ from $K_L^0 \rightarrow e^+e^-e^+e^-$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.01 ± 0.11 OUR AVERAGE</b>				Error includes scale factor of 1.6.
+0.13 ± 0.10 ± 0.03	200	LAI	05B NA48	
-0.09 ± 0.09 ± 0.02	441	ALAVI-HARATI01D	KTEV	$M_{ee} > 8 \text{ MeV}/c^2$

## CHARGE ASYMMETRY IN $\pi^+\pi^-\pi^0$ DECAYS

These are *CP*-violating charge-asymmetry parameters, defined at beginning of section "LINEAR COEFFICIENT  $g$  FOR  $K_L^0 \rightarrow \pi^+\pi^-\pi^0$  above.

See also note on Dalitz plot parameters in  $K^\pm$  section and note on "*CP* violation in  $K_L$  decays" above.

## LINEAR COEFFICIENT $j$ FOR $K_L^0 \rightarrow \pi^+\pi^-\pi^0$

VALUE	EVTS	DOCUMENT ID	TECN
<b>0.0012 ± 0.0008 OUR AVERAGE</b>			
0.0010 ± 0.0024 ± 0.0030	500k	ANGELOPO...	98c CPLR
-0.001 ± 0.011	6499	CHO	77
0.001 ± 0.003	4709	PEACH	77
0.0013 ± 0.0009	3M	SCRIBANO	70
0.0 ± 0.017	4400	SMITH	70 OSPK
0.001 ± 0.004	238k	BLANPIED	68

## QUADRATIC COEFFICIENT $f$ FOR $K_L^0 \rightarrow \pi^+\pi^-\pi^0$

VALUE	EVTS	DOCUMENT ID	TECN
<b>0.0045 ± 0.0024 ± 0.0059</b>	500k	ANGELOPO...	98c CPLR

## PARAMETERS for $K_L^0 \rightarrow \pi^+\pi^-\gamma$ DECAY

$$|\eta_{+-\gamma}| = |A(K_L^0 \rightarrow \pi^+\pi^-\gamma, \text{CP violating})/A(K_S^0 \rightarrow \pi^+\pi^-\gamma)|$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN
<b>2.35 ± 0.07 OUR AVERAGE</b>			
2.359 ± 0.062 ± 0.040	9045	MATTHEWS	95 E773
2.15 ± 0.26 ± 0.20	3671	RAMBERG	93B E731

## $\phi_{+-\gamma} = \text{phase of } \eta_{+-\gamma}$

VALUE (°)	EVTS	DOCUMENT ID	TECN
<b>44 ± 4 OUR AVERAGE</b>			
43.8 ± 3.5 ± 1.9	9045	MATTHEWS	95 E773
72 ± 23 ± 17	3671	RAMBERG	93B E731

## $|\epsilon'_{+-\gamma}|/\epsilon$ for $K_L^0 \rightarrow \pi^+\pi^-\gamma$

VALUE	CL%	EVTS	DOCUMENT ID	TECN
<b>&lt;0.3</b>	90	3671	<sup>1</sup> RAMBERG	93B E731

<sup>1</sup> RAMBERG 93B limit on  $|\epsilon'_{+-\gamma}|/\epsilon$  assumes than any difference between  $\eta_{+-}$  and  $\eta_{+-\gamma}$  is due to direct *CP* violation.

## $|g_{E1}|$ for $K_L^0 \rightarrow \pi^+\pi^-\gamma$

This parameter is the amplitude of the direct emission of a *CP* violating E1 electric dipole photon.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.21</b>	90	111k	ABOUZAIID	06A	KTEV $E_\gamma^* > 20 \text{ MeV}$

## T VIOLATION TESTS IN $K_L^0$ DECAYS

### $\text{Im}(\xi)$ in $K_{\mu 3}^0$ DECAY (from transverse $\mu$ pol.)

Test of *T* reversal invariance.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.007 ± 0.026 OUR AVERAGE</b>					
0.009 ± 0.030	12M		MORSE	80	CNTR Polarization
0.35 ± 0.30	207k		<sup>1</sup> CLARK	77	SPEC POL, $t=0$
-0.085 ± 0.064	2.2M		<sup>2</sup> SANDWEISS	73	CNTR POL, $t=0$
-0.02 ± 0.08			LONGO	69	CNTR POL, $t=3.3$
-0.2 ± 0.6			ABRAMS	68B	OSP K Polarization
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.012 ± 0.026			SCHMIDT	79	CNTR Repl. by MORSE 80

<sup>1</sup> CLARK 77 value has additional  $\xi(0)$  dependence  $+0.21\text{Re}[\xi(0)]$ .

<sup>2</sup> SANDWEISS 73 value corrected from value quoted in their paper due to new value of  $\text{Re}(\xi)$ . See footnote 4 of SCHMIDT 79.

## CPT-INVARIANCE TESTS IN $K_L^0$ DECAYS

### PHASE DIFFERENCE $\phi_{00} - \phi_{+-}$

Test of *CPT*.

OUR FIT is described in the note on "*CP* violation in  $K_L$  decays" in the  $K_L^0$  Particle Listings.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>0.006 ± 0.014 OUR FIT</b>			Error includes scale factor of 1.7. Assuming <i>CPT</i>
<b>0.34 ± 0.32 OUR FIT</b>			Not assuming <i>CPT</i>
0.006 ± 0.008	<sup>1</sup> SUPERWEAK	16	Assuming <i>CPT</i>
-0.30 ± 0.88	<sup>2</sup> SCHWINGEN...	95	Combined E731, E773
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.30 ± 0.35	<sup>3</sup> ABOUZAIID	11	KTEV Not assuming <i>CPT</i>
0.39 ± 0.22 ± 0.45	<sup>4</sup> ALAVI-HARATI03		KTEV
0.62 ± 0.71 ± 0.75	SCHWINGEN...	95	E773
-1.6 ± 1.2	<sup>5</sup> GIBBONS	93	E731
0.2 ± 2.6 ± 1.2	<sup>6</sup> CAROSI	90	NA31
-0.3 ± 2.4 ± 1.2	KARLSSON	90	E731

<sup>1</sup> SUPERWEAK 16 is a fake experiment to constrain  $\phi_{00} - \phi_{+-}$  to a small value as described in the note "*CP* violation in  $K_L$  decays."

<sup>2</sup> This SCHWINGENHEUER 95 values is the combined result of SCHWINGENHEUER 95 and GIBBONS 93, accounting for correlated systematic errors.

<sup>3</sup> Not independent of other phase parameters reported in ABOUZAIID 11.

<sup>4</sup> ALAVI-HARATI 03 fit  $\text{Re}(\epsilon'/\epsilon)$ ,  $\text{Im}(\epsilon'/\epsilon)$ ,  $\Delta m$ ,  $\tau_S$ , and  $\phi_{+-}$  simultaneously, not assuming *CPT*. Phase difference is obtained from  $\phi_{00} - \phi_{+-} \approx -3\text{Im}(\epsilon'/\epsilon)$  for small  $|\epsilon'/\epsilon|$ . Superseded by ABOUZAIID 11.

<sup>5</sup> GIBBONS 93 give detailed dependence of systematic error on lifetime (see the section on the  $K_S^0$  mean life) and mass difference (see the section on  $m_{K_0^0} - m_{K_S^0}$ ).

<sup>6</sup> CAROSI 90 is excluded from the fit because it it is not independent of  $\phi_{+-}$  and  $\phi_{00}$  values.

### PHASE DIFFERENCE $\phi_{+-} - \phi_{SW}$

Test of *CPT*. The Superweak phase  $\phi_{SW} \equiv \tan^{-1}(2\Delta m/\Delta\Gamma)$  where  $\Delta m = m_{K_L^0} - m_{K_S^0}$  and  $\Delta\Gamma = \hbar(\tau_L - \tau_S)/(\tau_L\tau_S)$ .

VALUE (°)	DOCUMENT ID	TECN
<b>0.61 ± 0.62 ± 1.01</b>	<sup>1</sup> ALAVI-HARATI03	KTEV

<sup>1</sup> ALAVI-HARATI 03 fit is the same as their  $\phi_{+-}$ ,  $\tau_{K_S}$ ,  $\Delta m$  fit, except that the parameter  $\phi_{+-} - \phi_{SW}$  is used in place of  $\phi$ .

$$\text{Re}(\frac{2}{3}\eta_{+-} + \frac{1}{3}\eta_{00}) - \frac{A_f}{A_\pi}$$

Test of *CPT*

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>-3 ± 35</b>	<sup>1</sup> ALAVI-HARATI02	E799	Uses $A_L$ from $K_{e3}$ decays

<sup>1</sup> ALAVI-HARATI 02 uses PDG 00 values of  $\eta_{+-}$  and  $\eta_{00}$ .

Meson Particle Listings

$K_L^0$

$\Delta S = \Delta Q$  IN  $K^0$  DECAYS

The relative amount of  $\Delta S \neq \Delta Q$  component present is measured by the parameter  $x$ , defined as

$$x = A(\overline{K}^0 \rightarrow \pi^- \ell^+ \nu) / A(K^0 \rightarrow \pi^- \ell^+ \nu) .$$

We list  $\text{Re}\{x\}$  and  $\text{Im}\{x\}$  for  $K_{e3}$  and  $K_{\mu 3}$  combined.

$$x = A(\overline{K}^0 \rightarrow \pi^- \ell^+ \nu) / A(K^0 \rightarrow \pi^- \ell^+ \nu) = A(\Delta S = -\Delta Q) / A(\Delta S = \Delta Q)$$

REAL PART OF  $x$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.0018 \pm 0.0041 \pm 0.0045$		ANGELOPO... 98D	CPLR	$K_{e3}$ from $K^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.10 $+0.18$ $-0.19$	79	SMITH	75B	WIRE $\pi^- p \rightarrow K^0 \Lambda$
0.04 $\pm 0.03$	4724	NIEBERGALL	74	ASPK $K^+ p \rightarrow K^0 p \pi^+$
$-0.008 \pm 0.044$	1757	ACKLER	73	OSPK $K_{e3}$ from $K^0$
$-0.03 \pm 0.07$	1367	HART	73	OSPK $K_{e3}$ from $K^0 \Lambda$
$-0.070 \pm 0.036$	1079	MALLARY	73	OSPK $K_{e3}$ from $K^0 \Lambda X$
0.03 $\pm 0.06$	410	<sup>1</sup> BURGUN	72	HBC $K^+ p \rightarrow K^0 p \pi^+$
0.04 $+0.10$ $-0.13$	100	<sup>2</sup> GRAHAM	72	OSPK $K_{\mu 3}$ from $K^0 \Lambda$
$-0.05 \pm 0.09$	442	<sup>2</sup> GRAHAM	72	OSPK $\pi^- p \rightarrow K^0 \Lambda$
0.26 $+0.10$ $-0.14$	126	MANN	72	HBC $K^- p \rightarrow n \overline{K}^0$
$-0.13 \pm 0.11$	342	<sup>2</sup> MANTSCH	72	OSPK $K_{e3}$ from $K^0 \Lambda$
0.04 $+0.07$ $-0.08$	222	<sup>1</sup> BURGUN	71	HBC $K^+ p \rightarrow K^0 p \pi^+$
0.25 $+0.07$ $-0.09$	252	WEBBER	71	HBC $K^- p \rightarrow n \overline{K}^0$
0.12 $\pm 0.09$	215	<sup>3</sup> CHO	70	DBC $K^+ d \rightarrow K^0 p p$
$-0.020 \pm 0.025$		<sup>4</sup> BENNETT	69	CNTR Charge asym+ Cu regen.
0.09 $+0.14$ $-0.16$	686	LITTENBERG	69	OSPK $K^+ n \rightarrow K^0 p$
0.03 $\pm 0.03$		<sup>4</sup> BENNETT	68	CNTR
0.09 $+0.07$ $-0.09$	121	JAMES	68	HBC $\overline{p} p$
0.17 $+0.16$ $-0.35$	116	FELDMAN	67B	OSPK $\pi^- p \rightarrow K^0 \Lambda$
0.17 $\pm 0.10$	335	<sup>3</sup> HILL	67	DBC $K^+ d \rightarrow K^0 p p$
0.035 $+0.11$ $-0.13$	196	AUBERT	65	HLBC $K^+$ charge exch.
0.06 $+0.18$ $-0.44$	152	<sup>5</sup> BALDO-...	65	HLBC $K^+$ charge exch.
$-0.08 \pm 0.16$ $-0.28$	109	<sup>6</sup> FRANZINI	65	HBC $\overline{p} p$

<sup>1</sup>BURGUN 72 is a final result which includes BURGUN 71.  
<sup>2</sup>First GRAHAM 72 value is second GRAHAM 72 value combined with MANTSCH 72.  
<sup>3</sup>CHO 70 is analysis of unambiguous events in new data and HILL 67.  
<sup>4</sup>BENNETT 69 is a reanalysis of BENNETT 68.  
<sup>5</sup>BALDO-CEOLIN 65 gives  $x$  and  $\theta$  converted by us to  $\text{Re}(x)$  and  $\text{Im}(x)$ .  
<sup>6</sup>FRANZINI 65 gives  $x$  and  $\theta$  for  $\text{Re}(x)$  and  $\text{Im}(x)$ . See SCHMIDT 67.

IMAGINARY PART OF  $x$

Assumes  $m_{K_L^0} - m_{K_S^0}$  positive. See Listings above.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.0012 \pm 0.0019 \pm 0.0009$	640k	ANGELOPO... 01B	CPLR	$K_{e3}$ from $K^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.0012 \pm 0.0019$	640k	<sup>1</sup> ANGELOPO...	98E	CPLR $K_{e3}$ from $K^0$
$-0.10 \pm 0.16$ $-0.19$	79	SMITH	75B	WIRE $\pi^- p \rightarrow K^0 \Lambda$
$-0.06 \pm 0.05$	4724	NIEBERGALL	74	ASPK $K^+ p \rightarrow K^0 p \pi^+$
$-0.017 \pm 0.060$	1757	ACKLER	73	OSPK $K_{e3}$ from $K^0$
0.09 $\pm 0.07$	1367	HART	73	OSPK $K_{e3}$ from $K^0 \Lambda$
0.107 $+0.092$ $-0.074$	1079	MALLARY	73	OSPK $K_{e3}$ from $K^0 \Lambda X$
0.07 $+0.06$ $-0.07$	410	<sup>2</sup> BURGUN	72	HBC $K^+ p \rightarrow K^0 p \pi^+$
0.12 $+0.17$ $-0.16$	100	<sup>3</sup> GRAHAM	72	OSPK $K_{\mu 3}$ from $K^0 \Lambda$
0.05 $\pm 0.13$	442	<sup>3</sup> GRAHAM	72	OSPK $\pi^- p \rightarrow K^0 \Lambda$
0.21 $+0.15$ $-0.12$	126	MANN	72	HBC $K^- p \rightarrow n \overline{K}^0$
$-0.04 \pm 0.16$	342	<sup>3</sup> MANTSCH	72	OSPK $K_{e3}$ from $K^0 \Lambda$
0.12 $+0.08$ $-0.09$	222	<sup>2</sup> BURGUN	71	HBC $K^+ p \rightarrow K^0 p \pi^+$
0.0 $\pm 0.08$	252	WEBBER	71	HBC $K^- p \rightarrow n \overline{K}^0$
$-0.08 \pm 0.07$	215	<sup>4</sup> CHO	70	DBC $K^+ d \rightarrow K^0 p p$
$-0.11 \pm 0.10$ $-0.11$	686	LITTENBERG	69	OSPK $K^+ n \rightarrow K^0 p$
$+0.22 \pm 0.37$ $-0.29$	121	JAMES	68	HBC $\overline{p} p$
0.0 $\pm 0.25$	116	FELDMAN	67B	OSPK $\pi^- p \rightarrow K^0 \Lambda$
$-0.20 \pm 0.10$	335	<sup>4</sup> HILL	67	DBC $K^+ d \rightarrow K^0 p p$
$-0.21 \pm 0.11$ $-0.15$	196	AUBERT	65	HLBC $K^+$ charge exch.
$-0.44 \pm 0.32$ $-0.19$	152	<sup>5</sup> BALDO-...	65	HLBC $K^+$ charge exch.
$+0.24 \pm 0.40$ $-0.30$	109	<sup>6</sup> FRANZINI	65	HBC $\overline{p} p$

<sup>1</sup>Superseded by ANGELOPOULOS 01B.  
<sup>2</sup>BURGUN 72 is a final result which includes BURGUN 71.  
<sup>3</sup>First GRAHAM 72 value is second GRAHAM 72 value combined with MANTSCH 72.  
<sup>4</sup>Footnote 10 of HILL 67 should read +0.58, not -0.58 (private communication) CHO 70 is analysis of unambiguous events in new data and HILL 67.  
<sup>5</sup>BALDO-CEOLIN 65 gives  $x$  and  $\theta$  converted by us to  $\text{Re}(x)$  and  $\text{Im}(x)$ .  
<sup>6</sup>FRANZINI 65 gives  $x$  and  $\theta$  for  $\text{Re}(x)$  and  $\text{Im}(x)$ . See SCHMIDT 67.

$K_L^0$  REFERENCES

BRFIT	16	RPP 2016 edition	C.-J. Lin	(PDG Collab.)
ETAFIT	16	RPP 2016 edition	C.-J. Lin	(PDG Collab.)
SUPERWEAK	16	RPP 2016 edition	C.-J. Lin	(PDG Collab.)
ABOUZAID	11	PR D83 092001	E. Abouzaid et al.	(FNAL KTeV Collab.)
ABOUZAID	11A	PRL 107 201803	E. Abouzaid et al.	(KTeV Collab.)
OGATA	11	PR D84 052009	R. Ogata et al.	(KEK E391a Collab.)
TUNG	11	PR D83 031101	Y.C. Tung et al.	(KEK E391a Collab.)
ABOUZAID	10	PR D81 052001	E. Abouzaid et al.	(FNAL KTeV Collab.)
AHN	10	PR D81 072004	J.K. Ahn et al.	(KEK E391a Collab.)
ABOUZAID	08	PR D77 112004	E. Abouzaid et al.	(FNAL KTeV Collab.)
ABOUZAID	08A	PR D78 032009	E. Abouzaid et al.	(FNAL KTeV Collab.)
ABOUZAID	08B	PR D78 032014	E. Abouzaid et al.	(FNAL KTeV Collab.)
ABOUZAID	08C	PRL 100 131803	E. Abouzaid et al.	(FNAL KTeV Collab.)
AHN	08	PRL 100 201802	J.K. Ahn et al.	(KEK E391a Collab.)
AMBROSINO	08F	EPJ C55 539	F. Ambrosino et al.	(KLOE Collab.)
ABOUZAID	07B	PRL 99 051804	E. Abouzaid et al.	(FNAL KTeV Collab.)
ABOUZAID	07C	PRL 99 081803	E. Abouzaid et al.	(FNAL KTeV Collab.)
ABOUZAID	07D	PR D76 052001	E. Abouzaid et al.	(FNAL KTeV Collab.)
AMBROSINO	07C	JHEP 0712 105	F. Ambrosino et al.	(KLOE Collab.)
ANDRE	07	ANP 322 2518	T. Andre	(EFI)
LAI	07	PL B645 26	A. Lai et al.	(CERN NA48 Collab.)
LAI	07A	PL B647 341	A. Lai et al.	(CERN NA48 Collab.)
NIX	07	PR D76 011101	J. Nix et al.	(KEK E391a Collab.)
ABOUZAID	06	PRL 96 101801	E. Abouzaid et al.	(KTeV Collab.)
ABOUZAID	06A	PR D74 032004	E. Abouzaid et al.	(KTeV Collab.)
Also		PR D74 039905 (errat.)	E. Abouzaid et al.	(KTeV Collab.)
ABOUZAID	06C	PR D74 097101	E. Abouzaid et al.	(KTeV Collab.)
AHN	06	PR D74 051105	J.K. Ahn et al.	(KEK E391a Collab.)
Also		PR D74 079901 (errat.)	J.K. Ahn et al.	(KEK E391a Collab.)
AMBROSINO	06D	PL B632 43	F. Ambrosino et al.	(KLOE Collab.)
AMBROSINO	05D	PL B636 166	F. Ambrosino et al.	(KLOE Collab.)
AMBROSINO	05F	PL B638 140	F. Ambrosino et al.	(KLOE Collab.)
BATLEY	06B	PL B633 173	J.R. Batley et al.	(CERN NA48/2 Collab.)
HILL	06	PR D74 096006	R.J. Hill	(FNAL)
PDG	06	JP G33 1	W.-M. Yao et al.	(PDG Collab.)
ALEXOPOU...	05	PR D71 012001	T. Alexopoulos et al.	(FNAL KTeV Collab.)
AMBROSINO	05C	PL B626 15	F. Ambrosino et al.	(KLOE Collab.)
CABIBBO	05	JHEP 0503 021	N. Cabibbo, G. Isidori	(CERN, ROMA1, FRAS)
LAI	05	PL B605 247	A. Lai et al.	(CERN NA48 Collab.)
LAI	05B	PL B615 31	A. Lai et al.	(CERN NA48 Collab.)
PARK	05	PRL 94 021801	H.K. Park et al.	(FNAL HyperCP Collab.)
ALAVI-HARATI	04A	PRL 93 021805	A. Alavi-Harati et al.	(FNAL KTeV/E799 Collab.)
ALEXOPOU...	04	PR D70 092006	T. Alexopoulos et al.	(FNAL KTeV Collab.)
ALEXOPOU...	04A	PR D70 092007	T. Alexopoulos et al.	(FNAL KTeV Collab.)
BATLEY	04	PL B595 75	J.R. Batley et al.	(CERN NA48 Collab.)
CIRIGLIANO	04	EPJ C35 53	V. Cirigliano, H. Neufeld, H. Pichl	(CIT, VALE+)
LAI	04B	PL B602 41	A. Lai et al.	(CERN NA48 Collab.)
LAI	04C	PL B604 1	A. Lai et al.	(CERN NA48 Collab.)
PDG	04	PL B592 1	S. Edelman et al.	(PDG Collab.)
SOZZI	04	EPJ C36 37	M. Sozzi	(PISA)
ADINOLFI	03	PL B566 61	M. Adinolfi et al.	(KLOE Collab.)
ALAVI-HARATI	03	PR D67 012005	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
Also		PR D70 079904 (errat.)	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	03B	PRL 90 141801	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
LAI	03	PL B551 7	A. Lai et al.	(CERN NA48 Collab.)
LAI	03C	EPJ C30 33	A. Lai et al.	(CERN NA48 Collab.)
ALAVI-HARATI	02	PRL 88 181601	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	02C	PRL 89 211801	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
BATLEY	02	PL B544 97	J.R. Batley et al.	(CERN NA48 Collab.)
CIRIGLIANO	02	EPJ C23 121	V. Cirigliano et al.	(VIEN, VALE, MARS)
LAI	02B	PL B536 229	A. Lai et al.	(CERN NA48 Collab.)
ALAVI-HARATI	01	PL B515 261	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01B	PRL 86 7614	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01D	PRL 86 5425	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01E	PRL 87 021801	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01F	PR D64 012003	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01G	PRL 87 071801	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01H	PRL 87 111802	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	01J	PR D64 112004	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ANGELOPO...	01	PL B503 49	A. Angelopoulos et al.	(CLEAR Collab.)
ANGELOPO...	01B	EPJ C22 95	A. Angelopoulos et al.	(CLEAR Collab.)
LAI	01B	PL B515 261	A. Lai et al.	(CERN NA48 Collab.)
LAI	01C	EPJ C22 231	A. Lai et al.	(CERN NA48 Collab.)
ALAVI-HARATI	00	PR D61 072006	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	00B	PRL 84 408	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	00D	PRL 84 5279	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	00E	PR D62 112001	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
AMBROSE	00	PRL 84 1389	D. Ambrose et al.	(BNL E871 Collab.)
APOSTOLA...	00	PL B473 186	A. Apostolakis et al.	(CLEAR Collab.)
PDG	00	EPJ C15 1	D.E. Groom et al.	(PDG Collab.)
ADAMS	99	PL B447 240	J. Adams et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	99B	PRL 83 917	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
ALAVI-HARATI	99D	PRL 83 22	A. Alavi-Harati et al.	(FNAL KTeV Collab.)
APOSTOLA...	99C	PL B458 545	A. Apostolakis et al.	(CLEAR Collab.)
Also		EPJ C18 41	A. Apostolakis et al.	(CLEAR Collab.)
FANTI	99B	PL B458 553	V. Fanti et al.	(CERN NA48 Collab.)
FANTI	99C	PL B465 335	V. Fanti et al.	(CERN NA48 Collab.)
MURAKAMI	99	PL B463 333	K. Murakami et al.	(KEK E162 Collab.)
ADAMS	98	PRL 80 4123	J. Adams et al.	(FNAL KTeV Collab.)
AMBROSE	98	PRL 81 4309	D. Ambrose et al.	(BNL E871 Collab.)
AMBROSE	98B	PRL 81 5734	D. Ambrose et al.	(BNL E871 Collab.)
ANGELOPO...	98	PL B420 191	A. Angelopoulos et al.	(CLEAR Collab.)
ANGELOPO...	98C	EPJ C5 389	A. Angelopoulos et al.	(CLEAR Collab.)
ANGELOPO...	98D	PL B444 38	A. Angelopoulos et al.	(CLEAR Collab.)
Also		EPJ C22 55	A. Angelopoulos et al.	(CLEAR Collab.)
ANGELOPO...	98E	PL B444 43	A. Angelopoulos et al.	(CLEAR Collab.)
ARISAKA	98	PL B432 230	K. Arisaka et al.	(FNAL E799 Collab.)
BENDER	98	PL B418 411	M. Bender et al.	(CERN NA48 Collab.)
DAMBROSIO	98	PL B423 385	G. D'Ambrosio, G. Isidori, J. Portoles	
SETZU	98	PL B420 205	M.G. Setzu et al.	
TAKEUCHI	98	PL B443 409	Y. Takeuchi et al.	(KYOT, KEK, HIRO)
FANTI	97	ZPHY C76 653	V. Fanti et al.	(CERN NA48 Collab.)
NOMURA	97	PL B408 445	T. Nomura et al.	(KYOT, KEK, HIRO)
ADLER	96B	ZPHY C70 211	R. Adler et al.	(CLEAR Collab.)
ADLER	96C	PL B369 367	R. Adler et al.	(CLEAR Collab.)
GU	96	PRL 76 4312	P. Gu et al.	(RUTG, UCLA, EFI, COLD+)
LEBER	96	PL B369 69	F. Leber et al.	(MANZ, CERN, EDIN, ORSAY+)
PDG	96	PR D54 1	R. M. Barnett et al.	(PDG Collab.)

ADLER	95	PL B363 237	R. Adler <i>et al.</i>	(CPLAR Collab.)	MCCARTHY	73	PR D7 687	R.L. McCarthy <i>et al.</i>	(LBL)
ADLER	95B	PL B363 243	R. Adler <i>et al.</i>	(CPLAR Collab.)	Also		PL 42B 291	R.L. McCarthy <i>et al.</i>	(LBL)
AKAGI	95	PR D51 2061	T. Akagi <i>et al.</i>	(TOHOK, TOKY, KYOT, KEK)	Also		Thesis LBL-550	R.L. McCarthy	(LBL)
BARR	95	ZPHY C65 361	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	MESSNER	73	PRL 30 876	R. Messner <i>et al.</i>	(COLO, SLAC, UCSC)
BARR	95C	PL B358 399	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	SANDWEISS	73	PRL 30 1002	J. Sandweiss <i>et al.</i>	(YALE, ANL)
HEINSON	95	PR D51 985	A.P. Heinson <i>et al.</i>	(BNL E791 Collab.)	WILLIAMS	73	PRL 31 1521	H.H. Williams <i>et al.</i>	(BNL, YALE)
KREUTZ	95	ZPHY C65 67	A. Kreutz <i>et al.</i>	(SIEG, EDIN, MANZ, ORSAY+)	ASHFORD	72	PL 38B 47	V.A. Ashford <i>et al.</i>	(UCSD)
MATTHEWS	95	PRL 75 2803	J.N. Matthews <i>et al.</i>	(RUTG, EFI, ELMT+)	BANNER	72B	PRL 29 237	M. Banner <i>et al.</i>	(PRIN)
SCHWINGEN...	95	PL 74 4376	B. Schwinger <i>et al.</i>	(EFI, CHIC+)	BARMIN	72B	SJNP 15 638	V.V. Barmin <i>et al.</i>	(ITEP)
SPENCER	95	PRL 74 3323	M.B. Spencer <i>et al.</i>	(UCLA, EFI, COLO+)			Translated from YAF 15 1152		
BARR	94	PL B328 528	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	BURGUN	72	NP B50 194	G. Burgun <i>et al.</i>	(SACL, CERN, OSLO)
GU	94	PRL 72 3000	P. Gu <i>et al.</i>	(RUTG, UCLA, EFI, COLO+)	GRAHAM	72	NC 9A 166	M.F. Graham <i>et al.</i>	(ILL, NEAS)
NAKAYA	94	PRL 73 2169	T. Nakaya <i>et al.</i>	(OSAK, UCLA, EFI, COLU+)	JAMES	72	NP B49 1	F. James <i>et al.</i>	(CERN, SACL, OSLO)
ROBERTS	94	PR D50 1874	D. Roberts <i>et al.</i>	(UCLA, EFI, COLU+)	KRENZ	72	LNC 4 213	W. Krenz <i>et al.</i>	(AACH, CERN, EDIN)
WEAVER	94	PRL 72 3758	M. Weaver <i>et al.</i>	(UCLA, EFI, COLU, ELMT+)	MANN	72	PR D6 137	W.A. Mann <i>et al.</i>	(MASA, BNL, YALE)
AKAGI	93	PR D47 R2644	T. Akagi <i>et al.</i>	(TOHOK, TOKY, KYOT, KEK)	MANTSCH	72	NC 9A 160	P.M. Mantsch <i>et al.</i>	(ILL, NEAS)
ARISAKA	93	PRL 70 1049	K. Arisaka <i>et al.</i>	(BNL E791 Collab.)	MCCARTHY	72	PL 42B 291	R.L. McCarthy <i>et al.</i>	(LBL)
ARISAKA	93B	PRL 71 3910	K. Arisaka <i>et al.</i>	(BNL E791 Collab.)	PICCONI	72	PRL 29 1412	R. Piccioni <i>et al.</i>	(SLAC)
BARR	93D	PL B317 233	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	Also		PR D9 2939	R. Piccioni <i>et al.</i>	(SLAC, UCSC, COLO)
GIBBONS	93	PRL 70 1199	L.K. Gibbons <i>et al.</i>	(FNAL E731 Collab.)	VOSBURGH	72	PR D6 1834	K.G. Vosburgh <i>et al.</i>	(RUTG, MASA)
Also		PR D55 6625	L.K. Gibbons <i>et al.</i>	(FNAL E731 Collab.)	Also		PR 26 866	K.G. Vosburgh <i>et al.</i>	(RUTG, MASA)
GIBBONS	93B	PRL 70 1203	L.K. Gibbons <i>et al.</i>	(FNAL E731 Collab.)	BALATS	71	SJNP 13 53	M.Y. Balats <i>et al.</i>	(ITEP)
GIBBONS	93C	Thesis RX-1487	L.K. Gibbons	(CHIC)			Translated from YAF 13 93		
Also		PR D55 6625	L.K. Gibbons <i>et al.</i>	(FNAL E731 Collab.)	BARMIN	71	PL 35B 604	V.V. Barmin <i>et al.</i>	(ITEP)
HARRIS	93	PRL 71 3914	D.A. Harris <i>et al.</i>	(EFI, UCLA, COLO+)	BURGUN	71	LNC 2 1169	G. Burgun <i>et al.</i>	(SACL, CERN, OSLO)
HARRIS	93B	PRL 71 3918	D.A. Harris <i>et al.</i>	(EFI, UCLA, COLO+)	CARNEGIE	71	PR D4 1	R.K. Carnegie <i>et al.</i>	(PRIN)
MAKOFF	93	PRL 70 1591	G. Makoff <i>et al.</i>	(FNAL E731 Collab.)	CHAN	71	Thesis LBL-350	J.H.S. Chan	(CMU, BNL, CASE)
Also		PRL 75 2069 (erratum)	G. Makoff <i>et al.</i>		CHO	71	PR D3 1557	Y. Cho <i>et al.</i>	(SLAC, STAN)
RAMBERG	93	PRL 70 2525	E. Ramberg <i>et al.</i>	(FNAL E731 Collab.)	ENSTROM	71	PR D4 2629	J. Enstrom <i>et al.</i>	(SLAC, STAN)
RAMBERG	93B	PRL 70 2529	E.J. Ramberg <i>et al.</i>	(FNAL E731 Collab.)	Also		Thesis SLAC-0125	J.E. Enstrom	(CERN, SACL, OSLO)
VAGINS	93	PRL 71 35	M.R. Vagins <i>et al.</i>	(BNL E845 Collab.)	JAMES	71	PL 35B 265	F. James <i>et al.</i>	(MASA, BNL, YALE)
ADLER	92B	PL B286 180	R. Adler <i>et al.</i>	(CPLAR Collab.)	MEISNER	71	PR D3 59	G.W. Meisner <i>et al.</i>	(ORSAY, CERN)
Also		SJNP 95 840	R. Adler <i>et al.</i>	(CPLAR Collab.)	REPELLIN	71	PL 36B 603	J.P. Repellin <i>et al.</i>	(LRL)
BARR	92	PL B284 440	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	WEBBER	71	PR D3 64	B.R. Webber <i>et al.</i>	(LRL)
GRAHAM	92	PL B295 169	G.E. Graham <i>et al.</i>	(FNAL E731 Collab.)	Also		PRL 21 498	B.R. Webber <i>et al.</i>	(LRL)
MORSE	92	PR D45 36	W.M. Morse <i>et al.</i>	(BNL, YALE, VASS)	Also		Thesis UCRL 19226	B.R. Webber	(ORSAY, CERN)
PDG	92	PR D45 51	K. Hikasa <i>et al.</i>	(KEK, LBL, BOST+)	WOLFF	71	PL 36B 517	B. Wolff <i>et al.</i>	(MCHS, DARE)
SOMALWAR	92	PRL 68 2580	S.V. Somalwar <i>et al.</i>	(FNAL E731 Collab.)	ALBROW	70	PL 33B 516	M.G. Albrow <i>et al.</i>	(EFI, ILLC, SLAC)
AKAGI	91B	PRL 67 2618	T. Akagi <i>et al.</i>	(TOHOK, TOKY, KYOT, KEK)	ARONSON	70	PRL 25 1057	S.H. Aronson <i>et al.</i>	(ITEP, JINR)
BARR	91	PL B259 389	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	BARMIN	70	PL 33B 377	V.V. Barmin <i>et al.</i>	(SACL)
HEINSON	91	PR D44 R1	A.P. Heinson <i>et al.</i>	(UCI, UCLA, LANL+)	BASILE	70	PR D2 78	P. Basile <i>et al.</i>	(ROCH)
PAPADIMITRI...	91	PR D44 R573	V. Papadimitriou <i>et al.</i>	(FNAL E731 Collab.)	BECHERRAWY	70	PR D1 1452	T. Becherrawy	(SLAC, JHU, UCLA)
BARKER	90	PRL 61 2541	K.R. Barker <i>et al.</i>	(FNAL E731 Collab.)	BUCHANAN	70	PL 33B 623	C.D. Buchanan <i>et al.</i>	(CERN, ORSAY, EPOL)
Also		PRL 61 2661	L.K. Gibbons <i>et al.</i>	(FNAL E731 Collab.)	Also		Private Comm.	A.J. Cox	(CERN, ORSAY, EPOL)
BARR	90B	PL B240 283	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	BUDAGOV	70	PR D2 815	I.A. Budagov <i>et al.</i>	(CMU, BNL, CASE)
BARR	90C	PL B242 523	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	Also		PL 28B 215	I.A. Budagov <i>et al.</i>	(BNL, CMU)
CAROSI	90	PL B237 303	R. Carosi <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	CHO	70	PR D1 3031	Y. Cho <i>et al.</i>	(CERN)
KARLSSON	90	PRL 64 2976	M. Karlsson <i>et al.</i>	(FNAL E731 Collab.)	Also		PRL 19 668	D.G. Hill <i>et al.</i>	(AACH, CERN, TOR)
OHL	90	PRL 64 2755	K.E. Ohi <i>et al.</i>	(BNL E845 Collab.)	CHOLLET	70	PL 31B 658	J.C. Chollet <i>et al.</i>	(COLU, HARV, CERN)
OHL	90B	PRL 65 1407	K.E. Ohi <i>et al.</i>	(BNL E845 Collab.)	CULLEN	70	PL 32B 523	M. Cullen <i>et al.</i>	(COLU)
PATTERSON	90	PRL 64 1491	J.R. Patterson <i>et al.</i>	(FNAL E731 Collab.)	MARX	70	PL 32B 219	J. Marx <i>et al.</i>	(PISA, COLU, HARV)
INAGAKI	89	PR D40 1712	T. Inagaki <i>et al.</i>	(KEK, TOKY, KYOT)	SCRIBANO	70	PL 32B 224	A. Scribano <i>et al.</i>	(UMD, BNL)
MATHIAZHA...	89	PRL 63 2181	C. Mathiazhagan <i>et al.</i>	(UCI, UCLA, LANL+)	SMITH	70	PL 32B 133	C.S. Smith <i>et al.</i>	(LRL)
MATHIAZHA...	89B	PRL 63 2185	C. Mathiazhagan <i>et al.</i>	(UCI, UCLA, LANL+)	WEBBER	70	PR D1 1967	B.R. Webber <i>et al.</i>	(PRIN)
WAHL	89	CERN-EP/89-96	H. Wahl	(CERN)	Also		Thesis UCRL 19226	B.R. Webber	(PRIN)
BARR	88	PL B214 303	G.D. Barr <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	BANNER	69	PR 188 2033	M. Banner <i>et al.</i>	(PRIN)
BURKHARDT	88	PL B206 169	H. Burkhardt <i>et al.</i>	(CERN, EDIN, MANZ, LALO+)	Also		PRL 21 1103	M. Banner <i>et al.</i>	(PRIN)
JASTRZEM...	88	PRL 61 2300	E. Jastrzemski <i>et al.</i>	(BNL, YALE)	Also		PRL 21 1107	J.W. Cronin, J.K. Liu, J.E. Pilcher	(COLU, BNL)
WOODS	88	PRL 60 1695	M. Woods <i>et al.</i>	(FNAL E731 Collab.)	BENNETT	69	PL 29B 317	S. Bennett <i>et al.</i>	(AACH3, CERN, TOR)
BURKHARDT	87	PL B199 139	H. Burkhardt <i>et al.</i>	(CERN, EDIN, MANZ+)	FAISSNER	69	PL 30B 204	H. Faissner <i>et al.</i>	(UCSD)
ARONSON	86	PR D33 3180	S.H. Aronson <i>et al.</i>	(BNL, CHIC, STAN+)	LITTENBERG	69	PR D2 654	L. Littenberg <i>et al.</i>	(MICH, UCLA)
Also		PL 170B 132	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)	LONGO	69	PR 181 1808	M.J. Longo, K.K. Young, J.A. Helland	(LRL)
PDG	86C	PR 170B 132	M. Aguilar-Benitez <i>et al.</i>	(CHIC, SACL)	PACIOTTI	69	Thesis UCRL 19446	M.A. Paciotti	(COLU)
COUPAL	85	PRL 55 566	D.P. Coupal <i>et al.</i>	(ITEP)	SAAL	69	Thesis	H.J. Saal	(ILL)
BALATS	83	SJNP 38 556	M.Y. Balats <i>et al.</i>		ABRAMS	68B	PR 176 1603	R.J. Abrams <i>et al.</i>	(CERN, ORSAY)
BERGSTROM	83	PL 131B 229	L. Bergstrom, E. Masso, P. Singer	(CERN)	ARNOLD	68B	PL 28B 56	R.G. Arnold <i>et al.</i>	(SACL)
ARONSON	82	PRL 48 1078	S.H. Aronson <i>et al.</i>	(BNL, CHIC, STAN+)	BASILE	68B	PL 28B 58	P. Basile <i>et al.</i>	(COLU, CERN)
ARONSON	82B	PRL 48 1306	S.H. Aronson <i>et al.</i>	(BNL, CHIC, PURD)	BENNETT	68	PL 27B 244	S. Bennett <i>et al.</i>	(CASE, HARV, MCGI)
Also		PL 116B 73	E. Fischbach <i>et al.</i>	(PURD, BNL, CHIC)	BLANPIED	68	PRL 21 1650	W.A. Blanpied <i>et al.</i>	(CERN, ORSAY, IPNP)
Also		PR D28 476	S.H. Aronson <i>et al.</i>	(BNL, CHIC, PURD)	BOHM	68B	PL 27B 594	A. Bohm <i>et al.</i>	(CERN, ORSAY, EPOL)
Also		PR D28 495	S.H. Aronson <i>et al.</i>	(BNL, CHIC, PURD)	BUDAGOV	68	NC 57A 182	I.A. Budagov <i>et al.</i>	(CERN, ORSAY, EPOL)
PDG	82B	PL 111B 70	M. Roos <i>et al.</i>	(HELS, CIT, CERN)	Also		PL 28B 215	F. James, H. Briand	(IPNP, CERN)
BIRULEV	81	NP B182 1	V.K. Birulev <i>et al.</i>	(JINR)	JAMES	68	PR 188 2033	J.A. Helland, M.J. Longo, K.K. Young	(UCLA, MICH)
Also		SJNP 31 622	V.K. Birulev <i>et al.</i>	(JINR)	Also		JETP 26 20	L.A. Kulyukina <i>et al.</i>	(JINR)
Also		Translated from YAF 31 1204			KULYUKINA	68	Translated from ZETF 53 29		
CARROLL	80B	PRL 44 529	A.S. Carroll <i>et al.</i>	(BNL, ROCH)	KUNZ	68	Thesis PU-68-46	P.F. Kunz	(PRIN)
CARROLL	80C	PL 96B 407	A.S. Carroll <i>et al.</i>	(BNL, ROCH)	BENNETT	67	PRL 19 993	S. Bennett <i>et al.</i>	(COLU)
CHO	80	PR D22 2688	Y. Cho <i>et al.</i>	(ANL, CMU)	DEBOUARD	67	NC 52A 662	X. de Bouard <i>et al.</i>	(CERN)
MORSE	80	PR D21 1750	W.M. Morse <i>et al.</i>	(BNL, YALE)	Also		PL 15 58	X. de Bouard <i>et al.</i>	(CERN, ORSAY, IPNP)
CHRISTENS...	79	PL 43 1209	J.H. Christenson <i>et al.</i>	(NYU)	DEVIN	67	PL 18 54	T.J. Devlin <i>et al.</i>	(PRIN, UMD)
SCHMIDT	79	PRL 43 556	M.P. Schmidt <i>et al.</i>	(YALE, BNL)	Also		PR 169 1045	G.A. Sayer <i>et al.</i>	(UMD, PPA, SLAC)
HILL	78	PL 73B 483	D.G. Hill <i>et al.</i>	(BNL, SLAC, SBER)	DORFAN	67	PRL 19 987	D.E. Dorfman <i>et al.</i>	(PENN)
CHO	77	PR D15 587	Y. Cho <i>et al.</i>	(ANL, CMU)	FELDMAN	67B	PR 155 1611	L. Feldman <i>et al.</i>	(PRIN)
CLARK	77	PR D15 553	A.R. Clark <i>et al.</i>	(LBL)	FITCH	67	PR 164 1711	V.L. Fitch <i>et al.</i>	(MASB)
Also		Thesis LBL-4275	G. Shen	(LBL)	GINSBERG	67	PR 162 1570	E.S. Ginsberg	(BNL, CMU)
DEVOE	77	PR D16 565	R. Devoe <i>et al.</i>	(EFI, ANL)	HILL	67	PRL 19 668	D.G. Hill <i>et al.</i>	(BNL)
PEACH	77	NP B127 399	K.J. Peach <i>et al.</i>	(BGNA, EDIN, GLAS+)	HOPKINS	67	PRL 19 185	H.W.K. Hopkins, T.C. Bacon, F.R. Eisler	(ILL)
BIRULEV	76	SJNP 24 178	V.K. Birulev <i>et al.</i>	(JINR)	NEFKENS	67	PR 157 1233	B.M.K. Nefkens <i>et al.</i>	(COLU)
Also		Translated from YAF 24 340			SCHMIDT	67	Thesis Nevis 160	P. Schmidt	(EPOL, MILA, PADO, ORSAY)
COOMBES	76	PRL 37 249	R.W. Coombes <i>et al.</i>	(STAN, NYU)	BEHR	66	PL 22 540	L. Behr <i>et al.</i>	(YALE)
GJESDAL	76	NP B109 118	G. Gjesdal <i>et al.</i>	(CERN, HEIDH)	HAWINS	66	PL 21 238	C.J.B. Hawins	(CERN, WISC)
BALDO...	75	NC 25A 688	M. Baldo-Ceolin <i>et al.</i>	(PADO, WIS)	Also		PR 156 1444	J.A. Anderson <i>et al.</i>	(LRL, WISC)
BLUMENTHAL	75	PRL 34 164	R.B. Blumenthal <i>et al.</i>	(PENN, CHIC, TEMP)	ANDERSON	65B	PRL 14 475	P. Astbury <i>et al.</i>	(CERN, ZUR)
BUCHANAN	75	PR D11 457	C.D. Buchanan <i>et al.</i>	(UCLA, SLAC, JHU)	ASTBURY	65B	PL 18 175	B. Aubert <i>et al.</i>	(EPOL, ORSAY)
CARITHERS	75	PRL 34 1244	W.C.J. Carithers <i>et al.</i>	(COLU, NYU)	AUBERT	65	PL 17 59	J.P. Lowys <i>et al.</i>	(PADO)
SMITH	75B	Thesis UCSD unpub.	J.G. Smith	(UCSD)	Also		PL 24B 75	M. Baldo-Ceolin <i>et al.</i>	(COLU, RUTG)
BISI	74	PL 50B 504	V. Bisi, M.J. Ferrero	(TORI)	BALDO...	65	NC 38 684	P. Franzini <i>et al.</i>	(BNL, YALE)
DONALDSON	74	Thesis SLAC-0184	G. Donaldson	(SLAC)	FRANZINI	65	PR 140B 127	P. Guidoni <i>et al.</i>	(VAND+)
Also		PR D14 2839	G. Donaldson <i>et al.</i>	(SLAC, UCSC)	GUIDONI	65	Argonne Conf. 49	H.W.K. Hopkins, T.C. Bacon, F. Eisler	(YERE)
DONALDSON	74B	PR D9 2960	G. Donaldson <i>et al.</i>	(SLAC, UCSC)	HOPKINS	65	Argonne Conf. 67	A.S. Aleksanyan <i>et al.</i>	(LEBD, MPEI, YERE)
Also		PL 31 337	G. Donaldson <i>et al.</i>	(CERN, HEIDH)	ALEKSANYAN	64B	Dubna Conf. 2 102		
GEWENIGER	74	PL 48B 483	C. Geweniger <i>et al.</i>	(CERN, HEIDH)	Also		JETP 19 1019		
Also		Thesis CERN Int. 74-4	V. Luth	(CERN, HEIDH)	ANIKINA	64	Translated from ZETF 46 1504	M.K. Anikina <i>et al.</i>	(GEOR, JINR)
GEWENIGER	74B	PL 48B 487	C. Geweniger <i>et al.</i>	(CERN, HEIDH)	Also		Translated from ZETF 46 59		
Also		PL 52B 119	S. Gjesdal <i>et al.</i>	(CERN, HEIDH)	FITCH	61	NC 22 1160	V.L. Fitch, P.A. Piroue, R.B. Perkins	(PRIN+)
GEWENIGER	74C	PL 52B 108	C. Geweniger <i>et al.</i>	(CERN, HEIDH)	GOOD	61	PR 124 1223	R.H. Good <i>et al.</i>	(LRL)
GJESDAL	74	PL 52B 113	S. Gjesdal <i>et al.</i>	(CERN, HEIDH)					
MESSNER	74	PRL 33 1458	R. Messner <i>et al.</i>	(COLO, SLAC, UCSC)					
NIEBERGALL	74	PL 49B 103	F. Niebergall <i>et al.</i>	(CERN, ORSAY, VIEN)					
WILLIAMS	74	PRL 33 240	H.H. Williams <i>et al.</i>	(BNL, YALE)					
ALEXANDER	73B	NP B65 301	G. Alexander <i>et al.</i>	(TELA, HEID)					

Meson Particle Listings

$K_L^0, K_0^*(800)$

DIB	92	PR D46 2265	C.O. Dib, R.D. Peccei	(UCLA)
KLEINKNECHT	92	CNPP 20 281	K. Kleinknecht	(MANZ)
New Results on CP Violation in Decays of Neutral K Mesons.				
KLEINKNECHT	90	ZPHY C46 557	K. Kleinknecht	(MANZ)
PEACH	90	JP G16 131	K.J. Peach	(EDIN)
BRYMAN	89	UMP A4 79	D.A. Bryman	(TRIU)
"Rare Kaon Decays"				
KLEINKNECHT	76	ARNS 26 1	K. Kleinknecht	(DORT)
GINSBERG	73	PR D8 3887	E.S. Ginsberg, J. Smith	(MIT, STON)
GINSBERG	70	PR D1 229	E.S. Ginsberg	(HAIF)
HEUSSE	70	LNC 3 449	P. Heusse <i>et al.</i>	(ORSAY)
CRONIN	68C	Vienna Conf. 281	J.W. Cronin	(PRIN)
RUBBIA	67	PL 24B 531	C. Rubbia, J. Steinberger	(CERN, COLU)
Also		PL 23 167	C. Rubbia, J. Steinberger	(CERN, COLU)
Also		PL 20 207	C. Alft-Steinberger <i>et al.</i>	(CERN)
Also		PL 21 595	C. Alft-Steinberger <i>et al.</i>	(CERN)
AUERBACH	66	PR 149 1052	L.B. Auerbach <i>et al.</i>	(PENN)
Also		PRL 14 192	L.B. Auerbach <i>et al.</i>	(PENN)
FIRESTONE	66B	PRL 17 116	A. Firestone <i>et al.</i>	(YALE, BNL)
BEHR	65	Argonne Conf. 59	L. Behr <i>et al.</i>	(EPOL, MILA, PADO)
MESTVIRISHVILI	65	JINR P 2449	A.N. Mestvirishvili <i>et al.</i>	(JINR)
TRILLING	65B	UCRL 16473	G.N. Trilling	(LRL)
Updated from 1965 Argonne Conference, page 115.				
JOVANOVICH	63	BNL Conf. 42	J.V. Jovanovich <i>et al.</i>	(BNL, UMD)

$K_0^*(800)$   
OR  $\kappa$

$I(J^P) = \frac{1}{2}(0^+)$

OMITTED FROM SUMMARY TABLE

Needs confirmation. See the mini-review on scalar mesons under  $f_0(500)$  (see the index for the page number).

$K_0^*(800)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>682 ± 29</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.		
826 ± 49	$^{+49}_{-34}$	1338	<sup>1</sup> ABLIKIM	11B BES2 $J/\psi \rightarrow K_S^0 K_S^0 \pi^+ \pi^-$
849 ± 77	$^{+18}_{-14}$	1421	<sup>2,3</sup> ABLIKIM	10E BES2 $J/\psi \rightarrow K^\pm K_S^0 \pi^\mp \pi^0$
841 ± 30	$^{+81}_{-73}$	25k	<sup>4,5</sup> ABLIKIM	06C BES2 $J/\psi \rightarrow \overline{K}^*(892)^0 K^+ \pi^-$
658 ± 13			<sup>6</sup> DESCOTES-G..06	RVUE $\pi K \rightarrow \pi K$
797 ± 19	± 43	15k	<sup>7,8</sup> AITALA	02 E791 $D^+ \rightarrow K^- \pi^+ \pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
663 ± 8 ± 34			<sup>9</sup> BUGG	10 RVUE S-matrix pole
706.0 ± 1.8 ± 22.8	141k		<sup>10</sup> BONVICINI	08A CLEO $D^+ \rightarrow K^- \pi^+ \pi^+$
856 ± 17 ± 13	54k		<sup>11</sup> LINK	07B FOCS $D^+ \rightarrow K^- \pi^+ \pi^+$
750 ± 30 ± 55			<sup>12</sup> BUGG	06 RVUE
855 ± 15	0.6k		<sup>13</sup> CAWLFIELD	06A CLEO $D^0 \rightarrow K^+ K^- \pi^0$
694 ± 53			<sup>3,14</sup> ZHOU	06 RVUE $K \rho \rightarrow K^- \pi^+ n$
753 ± 52			<sup>15</sup> PELAEZ	04A RVUE $K \pi \rightarrow K \pi$
594 ± 79			<sup>14</sup> ZHENG	04 RVUE $K^- \rho \rightarrow K^- \pi^+ n$
722 ± 60			<sup>16</sup> BUGG	03 RVUE $11 K^- \rho \rightarrow K^- \pi^+ n$
905 ± 65 ± 30			<sup>17</sup> ISHIDA	97B RVUE $11 K^- \rho \rightarrow K^- \pi^+ n$

<sup>1</sup> The Breit-Wigner parameters from a fit with seven intermediate resonances. The S-matrix pole position is  $(764 \pm 63^{+71}_{-54}) - i(306 \pm 149^{+143}_{-85})$  MeV.

<sup>2</sup> From a fit including ten additional resonances and energy-independent Breit-Wigner width.

<sup>3</sup> S-matrix pole.

<sup>4</sup> S-matrix pole. GUO 06 in a chiral unitary approach report a mass of  $757 \pm 33$  MeV and a width of  $558 \pm 82$  MeV.

<sup>5</sup> A fit in the  $K_0^*(800) + K^*(892) + K^*(1410)$  model with mass and width of the  $K_0^*(800)$  from ABLIKIM 06c well describes the left slope of the  $K_S^0 \pi^-$  invariant mass spectrum in  $\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$  decay studied by EPIFANOV 07.

<sup>6</sup> S-matrix pole. Using Roy-Steiner equations (ROY 71) as well as unitarity, analyticity and crossing symmetry constraints.

<sup>7</sup> Not seen by KOPP 01 using 7070 events of  $D^0 \rightarrow K^- \pi^+ \pi^0$ . LINK 02e and LINK 05i show clear evidence for a constant non-resonant scalar amplitude rather than  $K_0^*(800)$  in their high statistics analysis of  $D^+ \rightarrow K^- \pi^+ \mu^+ \nu_\mu$ .

<sup>8</sup> AUBERT 07t does not find evidence for the charged  $K_0^*(800)$  using 11k events of  $D^0 \rightarrow K^- K^+ \pi^0$ .

<sup>9</sup> S-Matrix pole. Supersedes BUGG 06. Combined analysis of ASTON 88, ABLIKIM 06c, AITALA 06, and LINK 09 using an s-dependent width with couplings to  $K\pi$  and  $K\eta'$ , and the Adler zero near thresholds.

<sup>10</sup> T-matrix pole.

<sup>11</sup> A Breit-Wigner mass and width.

<sup>12</sup> S-matrix pole. Reanalysis of ASTON 88, AITALA 02, and ABLIKIM 06c using for the  $\kappa$  an s-dependent width with an Adler zero near threshold.

<sup>13</sup> Breit-Wigner parameters. A significant S-wave can be also modeled as a non-resonant contribution.

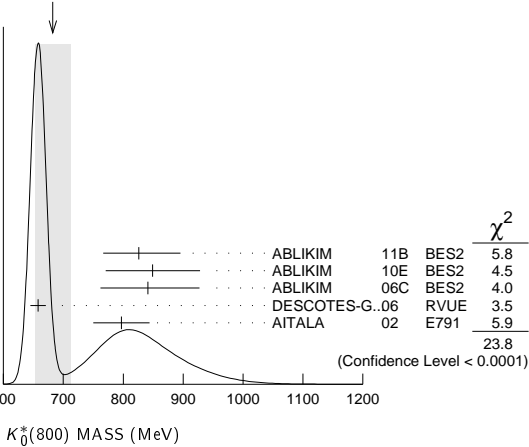
<sup>14</sup> Using ASTON 88.

<sup>15</sup> T-matrix pole. Reanalysis of data from LINGLIN 73, ESTABROOKS 78, and ASTON 88 in the unitarized ChPT model.

<sup>16</sup> T-matrix pole. Reanalysis of ASTON 88 data.

<sup>17</sup> Reanalysis of ASTON 88 using interfering Breit-Wigner amplitudes.

WEIGHTED AVERAGE  
682±29 (Error scaled by 2.4)



$K_0^*(800)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>547 ± 24</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.1.		
449 ± 156	$^{+144}_{-81}$	1338	<sup>18</sup> ABLIKIM	11B BES2 $J/\psi \rightarrow K_S^0 K_S^0 \pi^+ \pi^-$
512 ± 80	$^{+92}_{-44}$	1421	<sup>19,20</sup> ABLIKIM	10E BES2 $J/\psi \rightarrow K^\pm K_S^0 \pi^\mp \pi^0$
618 ± 90	$^{+96}_{-144}$	25k	<sup>19,21</sup> ABLIKIM	06C BES2 $J/\psi \rightarrow \overline{K}^*(892)^0 K^+ \pi^-$
557 ± 24			<sup>22</sup> DESCOTES-G..06	RVUE $\pi K \rightarrow \pi K$
410 ± 43 ± 87	15k	<sup>23,24</sup> AITALA	02 E791	$D^+ \rightarrow K^- \pi^+ \pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
658 ± 10 ± 44			<sup>25</sup> BUGG	10 RVUE S-matrix pole
638.8 ± 4.4 ± 40.4	141k		<sup>26</sup> BONVICINI	08A CLEO $D^+ \rightarrow K^- \pi^+ \pi^+$
464 ± 28 ± 22	54k		<sup>27</sup> LINK	07B FOCS $D^+ \rightarrow K^- \pi^+ \pi^+$
684 ± 120			<sup>28</sup> BUGG	06 RVUE
251 ± 48	0.6k		<sup>29</sup> CAWLFIELD	06A CLEO $D^0 \rightarrow K^+ K^- \pi^0$
606 ± 59			<sup>19,30</sup> ZHOU	06 RVUE $K \rho \rightarrow K^- \pi^+ n$
470 ± 66			<sup>31</sup> PELAEZ	04A RVUE $K \pi \rightarrow K \pi$
724 ± 332			<sup>30</sup> ZHENG	04 RVUE $K^- \rho \rightarrow K^- \pi^+ n$
772 ± 100			<sup>32</sup> BUGG	03 RVUE $11 K^- \rho \rightarrow K^- \pi^+ n$
545 ± 235 ± 110			<sup>33</sup> ISHIDA	97B RVUE $11 K^- \rho \rightarrow K^- \pi^+ n$

<sup>18</sup> The Breit-Wigner parameters from a fit with seven intermediate resonances. The S-matrix pole position is  $(764 \pm 63^{+71}_{-54}) - i(306 \pm 149^{+143}_{-85})$  MeV.

<sup>19</sup> S-matrix pole.

<sup>20</sup> From a fit including ten additional resonances and energy-independent Breit-Wigner width.

<sup>21</sup> A fit in the  $K_0^*(800) + K^*(892) + K^*(1410)$  model with mass and width of the  $K_0^*(800)$  from ABLIKIM 06c well describes the left slope of the  $K_S^0 \pi^-$  invariant mass spectrum in  $\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$  decay studied by EPIFANOV 07.

<sup>22</sup> S-matrix pole. Using Roy-Steiner equations (ROY 71) as well as unitarity, analyticity and crossing symmetry constraints.

<sup>23</sup> Not seen by KOPP 01 using 7070 events of  $D^0 \rightarrow K^- \pi^+ \pi^0$ . LINK 02e and LINK 05i show clear evidence for a constant non-resonant scalar amplitude rather than  $K_0^*(800)$  in their high statistics analysis of  $D^+ \rightarrow K^- \pi^+ \mu^+ \nu_\mu$ .

<sup>24</sup> AUBERT 07t does not find evidence for the charged  $K_0^*(800)$  using 11k events of  $D^0 \rightarrow K^- K^+ \pi^0$ .

<sup>25</sup> S-Matrix pole. Supersedes BUGG 06. Combined analysis of ASTON 88, ABLIKIM 06c, AITALA 06, and LINK 09 using an s-dependent width with couplings to  $K\pi$  and  $K\eta'$ , and the Adler zero near thresholds.

<sup>26</sup> T-matrix pole.

<sup>27</sup> A Breit-Wigner mass and width.

<sup>28</sup> S-matrix pole. Reanalysis of ASTON 88, AITALA 02, and ABLIKIM 06c using for the  $\kappa$  an s-dependent width with an Adler zero near threshold.

<sup>29</sup> Statistical error only. A fit to the Dalitz plot including the  $K_0^*(800)^\pm, K^*(892)^\pm$ , and  $\phi$  resonances modeled as Breit-Wigners. A significant S-wave can be also modeled as a non-resonant contribution.

<sup>30</sup> Using ASTON 88.

<sup>31</sup> T-matrix pole. Reanalysis of data from LINGLIN 73, ESTABROOKS 78, and ASTON 88 in the unitarized ChPT model.

<sup>32</sup> T-matrix pole. Reanalysis of ASTON 88 data.

<sup>33</sup> Reanalysis of ASTON 88 using interfering Breit-Wigner amplitudes.

$K_0^*(800)$  REFERENCES

ABLIKIM	11B	PL B698 183	M. Ablikim <i>et al.</i>	(BES II Collab.)
ABLIKIM	10E	PL B693 88	M. Ablikim <i>et al.</i>	(BES II Collab.)
BUGG	10	PR D81 014002	D.V. Bugg	(LOQM)
LINK	09	PL B681 14	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
BONVICINI	08A	PR D78 052001	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
AUBERT	07T	PR D76 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
EPIFANOV	07	PL B654 65	D. Epifanov <i>et al.</i>	(BELLE Collab.)
LINK	07B	PL B653 1	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)

See key on page 601

## Meson Particle Listings

 $K_0^*(800), K^*(892)$ 

ABLIKIM	06C	PL B633 681	M. Ablikim <i>et al.</i>	(BES Collab.)
AITALA	06	PR D73 032004	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
Also		PR D74 059901 (errat.)	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BUGG	06	PL B632 471	D.V. Bugg	(LOQM)
CRAWFIELD	06A	PR D74 031108	C. Crawford <i>et al.</i>	(CLEO Collab.)
DESCOTES-G...	06	EPJ C48 553	S. Descotes-Genon, B. Moussallam	
GUO	06	NP A773 78	F.K. Guo <i>et al.</i>	
ZHOU	06	NP A775 212	Z.Y. Zhou, H.Q. Zheng	
LINK	05I	PL B621 72	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
PELAEZ	05A	MPL A19 2879	J.R. Pelaez	
ZHENG	04	NP A733 235	H.Q. Zheng <i>et al.</i>	
BUGG	03	PL B572 1	D.V. Bugg	
AITALA	02	PRL 89 121801	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
LINK	02E	PL B535 43	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
KOPP	01	PR D63 092001	S. Kopp <i>et al.</i>	(CLEO Collab.)
ISHIDA	97B	PTP 98 621	S. Ishida <i>et al.</i>	
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ESTABROOKS	78	NP B133 490	P.G. Estabrooks <i>et al.</i>	(MCGI, CARL, DURH+)
LINGLIN	73	NP B95 408	D. Linglin	(CERN)
ROY	71	PL 36B 353	S.M. Roy	

 $K^*(892)$ 

$$I(J^P) = \frac{1}{2}(1^-)$$

 $K^*(892)$  MASS

## CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>891.66 ± 0.26 OUR AVERAGE</b>					
892.6 ± 0.5	5840	BAUBILLIER 84B	HBC	—	8.25 $K^- p \rightarrow \bar{K}^0 \pi^- p$
888 ± 3		NAPIER 84	SPEC	+	200 $\pi^- p \rightarrow 2K_S^0 X$
891 ± 1		NAPIER 84	SPEC	—	200 $\pi^- p \rightarrow 2K_S^0 X$
891.7 ± 2.1	3700	BARTH 83	HBC	+	70 $K^+ p \rightarrow K^0 \pi^+ X$
891 ± 1	4100	TOAFF 81	HBC	—	6.5 $K^- p \rightarrow \bar{K}^0 \pi^- p$
892.8 ± 1.6		AJINENKO 80	HBC	+	32 $K^+ p \rightarrow K^0 \pi^+ X$
890.7 ± 0.9	1800	AGUILAR... 78B	HBC	±	0.76 $\bar{p} p \rightarrow K^+ \bar{K}^0 \pi^\pm$
886.6 ± 2.4	1225	BALAND 78	HBC	±	12 $\bar{p} p \rightarrow (K\pi)^\pm X$
891.7 ± 0.6	6706	COOPER 78	HBC	±	0.76 $\bar{p} p \rightarrow (K\pi)^\pm X$
891.9 ± 0.7	9000	<sup>1</sup> PALER 75	HBC	—	14.3 $K^- p \rightarrow (K\pi)^- X$
892.2 ± 1.5	4404	AGUILAR... 71B	HBC	—	3.9, 4.6 $K^- p \rightarrow (K\pi)^- p$
891 ± 2	1000	CRENNELL 69D	DBC	—	3.9 $K^- N \rightarrow K^0 \pi^- X$
890 ± 3.0	720	BARLOW 67	HBC	±	1.2 $\bar{p} p \rightarrow (K^0 \pi)^\pm K^\mp$
889 ± 3.0	600	BARLOW 67	HBC	±	1.2 $\bar{p} p \rightarrow (K^0 \pi)^\pm K\pi$
891 ± 2.3	620	<sup>2</sup> DEBAERE 67B	HBC	+	3.5 $K^+ p \rightarrow K^0 \pi^+ p$
891.0 ± 1.2	1700	<sup>3</sup> WOJCICKI 64	HBC	—	1.7 $K^- p \rightarrow \bar{K}^0 \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
893.5 ± 1.1	27k	<sup>4</sup> ABELE 99D	CBAR	±	0.0 $\bar{p} p \rightarrow K^+ K^- \pi^0$
890.4 ± 0.2 ± 0.5	80 ± 0.8k	<sup>5</sup> BIRD 89	LASS	—	11 $K^- p \rightarrow \bar{K}^0 \pi^- p$
890.0 ± 2.3	800	<sup>2,3</sup> CLELAND 82	SPEC	+	30 $K^+ p \rightarrow K_S^0 \pi^+ p$
896.0 ± 1.1	3200	<sup>2,3</sup> CLELAND 82	SPEC	+	50 $K^+ p \rightarrow K_S^0 \pi^+ p$
893 ± 1	3600	<sup>2,3</sup> CLELAND 82	SPEC	—	50 $K^+ p \rightarrow K_S^0 \pi^- p$
896.0 ± 1.9	380	DELFOSSSE 81	SPEC	+	50 $K^\pm p \rightarrow K^\pm \pi^0 p$
886.0 ± 2.3	187	DELFOSSSE 81	SPEC	—	50 $K^\pm p \rightarrow K^\pm \pi^0 p$
894.2 ± 2.0	765	<sup>2</sup> CLARK 73	HBC	—	3.13 $K^- p \rightarrow \bar{K}^0 \pi^- p$
894.3 ± 1.5	1150	<sup>2,3</sup> CLARK 73	HBC	—	3.3 $K^- p \rightarrow \bar{K}^0 \pi^- p$
892.0 ± 2.6	341	<sup>2</sup> SCHWEING...68	HBC	—	5.5 $K^- p \rightarrow \bar{K}^0 \pi^- p$

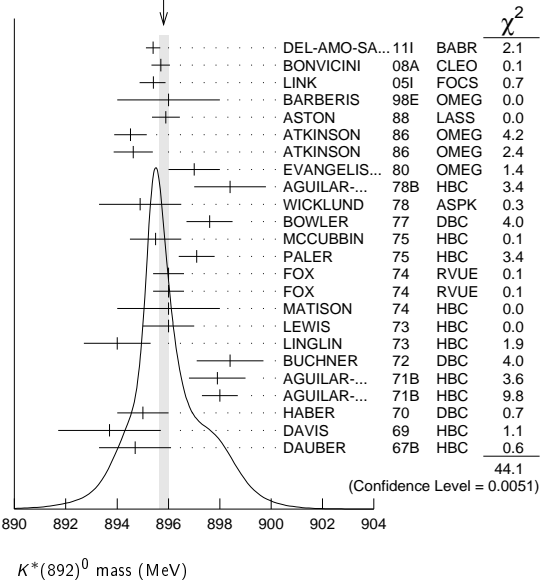
CHARGED ONLY, PRODUCED IN  $\tau$  LEPTON DECAYS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>895.47 ± 0.20 ± 0.74</b>	53k	<sup>6</sup> EPIFANOV 07	BELL	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
892.0 ± 0.5		<sup>7</sup> BOITO 10	RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
892.0 ± 0.9		<sup>8,9</sup> BOITO 09	RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
895.3 ± 0.2		<sup>8,10</sup> JAMIN 08	RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
896.4 ± 0.9	11970	<sup>11</sup> BONVICINI 02	CLEO	$\tau^- \rightarrow K^- \pi^0 \nu_\tau$
895 ± 2		<sup>12</sup> BARATE 99R	ALEP	$\tau^- \rightarrow K^- \pi^0 \nu_\tau$

## NEUTRAL ONLY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>895.81 ± 0.19 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
895.4 ± 0.2 ± 0.2	243k	<sup>13</sup> DEL-AMO-SA...11I	BABR	$D^+ \rightarrow K^- \pi^+ e^+ \nu_e$
895.7 ± 0.2 ± 0.3	141k	<sup>14</sup> BONVICINI 08A	CLEO	$D^+ \rightarrow K^- \pi^+ \pi^+$
895.41 ± 0.32 ± 0.35 -0.43	18k	<sup>15</sup> LINK 05A	FOCS	$D^+ \rightarrow K^- \pi^+ \mu^+ \nu_\mu$
896 ± 2		BARBERIS 98E	OMEG	450 $pp \rightarrow p_f p_S K^* \bar{K}^*$
895.9 ± 0.5 ± 0.2		ASTON 88	LASS	11 $K^- p \rightarrow K^- \pi^+ n$
894.52 ± 0.63	25k	<sup>1</sup> ATKINSON 86	OMEG	20-70 $\gamma p$
894.63 ± 0.76	20k	<sup>1</sup> ATKINSON 86	OMEG	20-70 $\gamma p$
897 ± 1	28k	EVANGELIS... 80	OMEG	10 $\pi^- p \rightarrow K^+ \pi^- (\Lambda, \Sigma)$
898.4 ± 1.4	1180	AGUILAR... 78B	HBC	0.76 $\bar{p} p \rightarrow K^\mp K_S^0 \pi^\pm$
894.9 ± 1.6		WICKLUND 78	ASPK	3.4, 6 $K^\pm N \rightarrow (K\pi)^0 N$
897.6 ± 0.9		BOWLER 77	DBC	5.4 $K^+ d \rightarrow K^+ \pi^- pp$
895.5 ± 1.0	3600	MCCUBBIN 75	HBC	3.6 $K^- p \rightarrow K^- \pi^+ n$

897.1 ± 0.7	22k	<sup>1</sup> PALER 75	HBC	14.3 $K^- p \rightarrow (K\pi)^0 X$
896.0 ± 0.6	10k	FOX 74	RVUE	2 $K^- p \rightarrow K^- \pi^+ n$
896.0 ± 0.6		FOX 74	RVUE	2 $K^+ n \rightarrow K^+ \pi^- p$
896 ± 2		<sup>16</sup> MATISON 74	HBC	12 $K^+ p \rightarrow K^+ \pi^- \Delta$
896 ± 1	3186	LEWIS 73	HBC	2.1-2.7 $K^+ p \rightarrow K\pi\pi p$
894.0 ± 1.3		<sup>16</sup> LINGLIN 73	HBC	2-13 $K^+ p \rightarrow K^+ \pi^- \pi^+ p$
898.4 ± 1.3	1700	<sup>2</sup> BUCHNER 72	DBC	4.6 $K^+ n \rightarrow K^+ \pi^- p$
897.9 ± 1.1	2934	<sup>2</sup> AGUILAR... 71B	HBC	3.9, 4.6 $K^- p \rightarrow K^- \pi^+ n$
898.0 ± 0.7	5362	<sup>2</sup> AGUILAR... 71B	HBC	3.9, 4.6 $K^- p \rightarrow K^- \pi^+ \pi^- p$
895 ± 1	4300	<sup>3</sup> HABER 70	DBC	3 $K^- N \rightarrow K^- \pi^+ X$
893.7 ± 2.0	10k	DAVIS 69	HBC	12 $K^+ p \rightarrow K^+ \pi^- \pi^+ p$
894.7 ± 1.4	1040	<sup>2</sup> DAUBER 67B	HBC	2.0 $K^- p \rightarrow K^- \pi^+ \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
895.53 ± 0.17		LEES 13F	BABR	$D^+ \rightarrow K^+ K^- \pi^+$
894.9 ± 0.5 ± 0.7	14.4k	<sup>17</sup> MITCHELL 09A	CLEO	$D^+ \rightarrow K^+ K^- \pi^+$
896.2 ± 0.3	20k	<sup>8</sup> AUBERT 07AK	BABR	10.6 $e^+ e^- \rightarrow K^{*0} K^\pm \pi^\mp \gamma$
900.7 ± 1.1	5900	BARTH 83	HBC	70 $K^+ p \rightarrow K^+ \pi^- X$

WEIGHTED AVERAGE  
895.81 ± 0.19 (Error scaled by 1.4)

- Inclusive reaction. Complicated background and phase-space effects.
- Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ . See note.
- Number of events in peak reevaluated by us.
- K-matrix pole.
- From a partial wave amplitude analysis.
- From a fit in the  $K_0^*(800) + K^*(892) + K^*(1410)$  model.
- From the pole position of the  $K\pi$  vector form factor using EPIFANOV 07 and constraints from  $K_{J3}$  decays in ANTONELLI 10.
- Systematic uncertainties not estimated.
- From the pole position of the  $K\pi$  vector form factor in the complex s-plane and using EPIFANOV 07 data.
- Reanalysis of EPIFANOV 07 using resonance chiral theory.
- Calculated by us from the shift by  $4.7 \pm 0.9$  MeV (statistical uncertainty only) reported in BONVICINI 02 with respect to the world average value from PDG 00.
- With mass and width of the  $K^*(1410)$  fixed at 1412 MeV and 227 MeV, respectively.
- Taking into account the  $K^*(892)^0$ , S-wave and P-wave ( $K^*(1410)^0$ ).
- From the isobar model with a complex pole for the  $\kappa$ .
- Fit to  $K\pi$  mass spectrum includes a non-resonant scalar component.
- From pole extrapolation.
- This value comes from a fit with  $\chi^2$  of 178/117.

 $K^*(892)$  MASSES AND MASS DIFFERENCES

Unrealistically small errors have been reported by some experiments. We use simple “realistic” tests for the minimum errors on the determination of a mass and width from a sample of  $N$  events:

$$\delta_{\min}(m) = \frac{\Gamma}{\sqrt{N}}, \quad \delta_{\min}(\Gamma) = 4 \frac{\Gamma}{\sqrt{N}}. \quad (1)$$

# Meson Particle Listings

## $K^*(892)$

We consistently increase unrealistic errors before averaging. For a detailed discussion, see the 1971 edition of this Note.

$m_{K^*(892)^0} - m_{K^*(892)^\pm}$					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>6.7±1.2 OUR AVERAGE</b>					
7.7±1.7	2980	AGUILAR-...	78B HBC	±0	0.76 $\overline{p}p \rightarrow K^\mp K_S^0 \pi^\pm$
5.7±1.7	7338	AGUILAR-...	71B HBC	-0	3.9,4.6 $K^-p$
6.3±4.1	283	18 BARASH	67B HBC		0.0 $\overline{p}p$
18 Number of events in peak reevaluated by us.					

### $K^*(892)$ RANGE PARAMETER

All from partial wave amplitude analyses.

VALUE (GeV <sup>-1</sup> )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
2.1 ± 0.5 ± 0.5	243k	19 DEL-AMO-SA.111	BABR	0	$D^+ \rightarrow K^- \pi^+ e^+ \nu_e$
3.96 ± 0.54 <sup>+1.31</sup> <sub>-0.90</sub>	18k	20 LINK	05I FOCUS	0	$D^+ \rightarrow K^- \pi^+ \mu^+ \nu_\mu$
3.4 ± 0.7		ASTON	88 LASS	0	11 $K^-p \rightarrow K^- \pi^+ n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
12.1 ± 3.2 ± 3.0		BIRD	89 LASS	-	11 $K^-p \rightarrow \overline{K}^0 \pi^- p$
19 Taking into account the $K^*(892)^0$ , S-wave and P-wave ( $K^*(1410)^0$ ).					
20 Fit to $K\pi$ mass spectrum includes a non-resonant scalar component.					

### $K^*(892)$ WIDTH

#### CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>50.8±0.9 OUR FIT</b>					
<b>50.8±0.9 OUR AVERAGE</b>					
49 ± 2	5840	BAUBILLIER	84B HBC	-	8.25 $K^-p \rightarrow \overline{K}^0 \pi^- p$
56 ± 4		NAPIER	84 SPEC	-	200 $\pi^-p \rightarrow 2K_S^0 X$
51 ± 2	4100	TOAFF	81 HBC	-	6.5 $K^-p \rightarrow \overline{K}^0 \pi^- p$
50.5±5.6		AJINENKO	80 HBC	+	32 $K^+p \rightarrow K^0 \pi^+ X$
45.8±3.6	1800	AGUILAR-...	78B HBC	±	0.76 $\overline{p}p \rightarrow K^\mp K_S^0 \pi^\pm$
52.0±2.5	6706	21 COOPER	78 HBC	±	0.76 $\overline{p}p \rightarrow (K\pi)^\pm X$
52.1±2.2	9000	22 PALER	75 HBC	-	14.3 $K^-p \rightarrow (K\pi)^-$
46.3±6.7	765	21 CLARK	73 HBC	-	3.13 $K^-p \rightarrow \overline{K}^0 \pi^- p$
48.2±5.7	1150	21,23 CLARK	73 HBC	-	3.3 $K^-p \rightarrow \overline{K}^0 \pi^- p$
54.3±3.3	4404	21 AGUILAR-...	71B HBC	-	3.9,4.6 $K^-p \rightarrow (K\pi)^- p$
46 ± 5	1700	21,23 WOJCICKI	64 HBC	-	1.7 $K^-p \rightarrow \overline{K}^0 \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
54.8±1.7	27k	24 ABELE	99D CBAR	±	0.0 $\overline{p}p \rightarrow K^+ K^- \pi^0$
45.2±1 ± 2	79.7±0.8k	25 BIRD	89 LASS	-	11 $K^-p \rightarrow \overline{K}^0 \pi^- p$
42.8±7.1	3700	BARTH	83 HBC	+	70 $K^+p \rightarrow K^0 \pi^+ X$
64.0±9.2	800	21,23 CLELAND	82 SPEC	+	30 $K^+p \rightarrow K_S^0 \pi^+ p$
62.0±4.4	3200	21,23 CLELAND	82 SPEC	+	50 $K^+p \rightarrow K_S^0 \pi^+ p$
55 ± 4	3600	21,23 CLELAND	82 SPEC	-	50 $K^+p \rightarrow K_S^0 \pi^- p$
62.6±3.8	380	DELFOSSSE	81 SPEC	+	50 $K^\pm p \rightarrow K^\pm \pi^0 p$
50.5±3.9	187	DELFOSSSE	81 SPEC	-	50 $K^\pm p \rightarrow K^\pm \pi^0 p$

#### CHARGED ONLY, PRODUCED IN $\tau$ LEPTON DECAYS

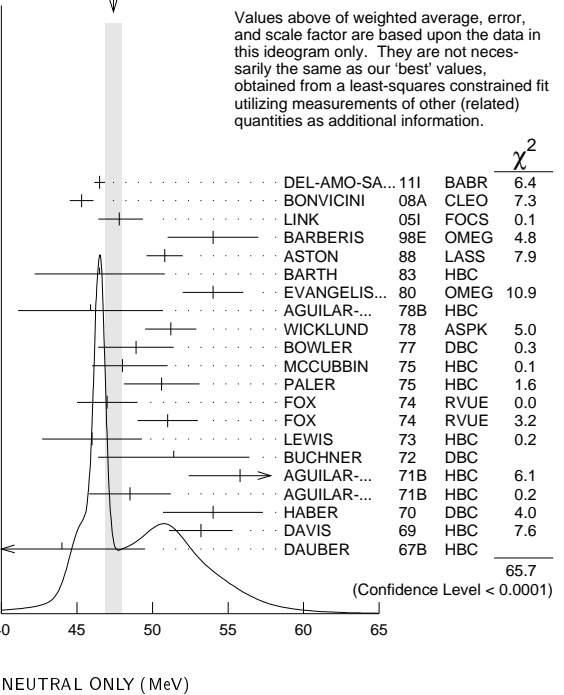
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>46.2±0.6±1.2</b>				
53k	26 EPIFANOV	07 BELL	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
46.5±1.1		27 BOITO	10 RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
46.2±0.4		28,29 BOITO	09 RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
47.5±0.4		28,30 JAMIN	08 RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
55 ± 8		31 BARATE	99R ALEP	$\tau^- \rightarrow K^- \pi^0 \nu_\tau$

#### NEUTRAL ONLY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>47.4 ± 0.6 OUR FIT</b>				
<b>47.4 ± 0.6 OUR AVERAGE</b>				
46.5 ± 0.3 ± 0.2	243k	32 DEL-AMO-SA...111	BABR	$D^+ \rightarrow K^- \pi^+ e^+ \nu_e$
45.3 ± 0.5 ± 0.6	141k	33 BONVICINI	08A CLEO	$D^+ \rightarrow K^- \pi^+ \pi^+$
47.79 ± 0.86 <sup>+1.32</sup> <sub>-1.06</sub>	18k	34 LINK	05I FOCUS	$D^+ \rightarrow K^- \pi^+ \mu^+ \nu_\mu$
54 ± 3		BARBERIS	98E OMEG	450 $pp \rightarrow p_f p_S K^* \overline{K}^*$
50.8 ± 0.8 ± 0.9		ASTON	88 LASS	11 $K^-p \rightarrow K^- \pi^+ n$
46.5 ± 4.3	5900	BARTH	83 HBC	70 $K^+p \rightarrow K^+ \pi^- X$
54 ± 2	28k	EVANGELIS...	80 OMEG	10 $\pi^-p \rightarrow K^+ \pi^- (\Lambda, \Sigma)$
45.9 ± 4.8	1180	AGUILAR-...	78B HBC	0.76 $\overline{p}p \rightarrow K^\mp K_S^0 \pi^\pm$
51.2 ± 1.7		WICKLUND	78 ASPK	3,4,6 $K^\pm N \rightarrow (K\pi)^0 N$
48.9 ± 2.5		BOWLER	77 DBC	5.4 $K^+d \rightarrow K^+ \pi^- pp$
48 <sup>+3</sup> <sub>-2</sub>	3600	MCCUBBIN	75 HBC	3.6 $K^-p \rightarrow K^- \pi^+ n$

50.6 ± 2.5	22k	22 PALER	75 HBC	14.3 $K^-p \rightarrow (K\pi)^0 X$
47 ± 2	10k	FOX	74 RVUE	2 $K^-p \rightarrow K^- \pi^+ n$
51 ± 2		FOX	74 RVUE	2 $K^+n \rightarrow K^+ \pi^- p$
46.0 ± 3.3	3186	21 LEWIS	73 HBC	2.1-2.7 $K^+p \rightarrow K\pi\pi p$
51.4 ± 5.0	1700	21 BUCHNER	72 DBC	4.6 $K^+n \rightarrow K^+ \pi^- p$
55.8 <sup>+4.2</sup> <sub>-3.4</sub>	2934	21 AGUILAR-...	71B HBC	3.9,4.6 $K^-p \rightarrow K^- \pi^+ n$
48.5 ± 2.7	5362	AGUILAR-...	71B HBC	3.9,4.6 $K^-p \rightarrow K^- \pi^+ \pi^- p$
54.0 ± 3.3	4300	21,23 HABER	70 DBC	3 $K^-N \rightarrow K^- \pi^+ X$
53.2 ± 2.1	10k	21 DAVIS	69 HBC	12 $K^+p \rightarrow K^+ \pi^- \pi^+ p$
44 ± 5.5	1040	21 DAUBER	67B HBC	2.0 $K^-p \rightarrow K^- \pi^+ \pi^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
44.90 ± 0.30		LEES	13F BABR	$D^+ \rightarrow K^+ K^- \pi^+$
45.7 ± 1.1 ± 0.5	14.4k	35 MITCHELL	09A CLEO	$D_S^+ \rightarrow K^+ K^- \pi^+$
50.6 ± 0.9	20k	28 AUBERT	07AK BABR	10.6 $e^+ e^- \rightarrow K^{*0} K^\pm \pi^\mp \gamma$

WEIGHTED AVERAGE  
47.4±0.6 (Error scaled by 2.0)



- 21 Width errors enlarged by us to  $4 \times \Gamma/\sqrt{N}$ ; see note.  
22 Inclusive reaction. Complicated background and phase-space effects.  
23 Number of events in peak reevaluated by us.  
24 K-matrix pole.  
25 From a partial wave amplitude analysis.  
26 From a fit in the  $K_S^0(800) + K^*(892) + K^*(1410)$  model.  
27 From the pole position of the  $K\pi$  vector form factor using EPIFANOV 07 and constraints from  $K_{J3}$  decays in ANTONELLI 10.  
28 Systematic uncertainties not estimated.  
29 From the pole position of the  $K\pi$  vector form factor in the complex s-plane and using EPIFANOV 07 data.  
30 Reanalysis of EPIFANOV 07 using resonance chiral theory.  
31 With mass and width of the  $K^*(1410)$  fixed at 1412 MeV and 227 MeV, respectively.  
32 Taking into account the  $K^*(892)^0$ , S-wave and P-wave ( $K^*(1410)^0$ ).  
33 From the isobar model with a complex pole for the  $\kappa$ .  
34 Fit to  $K\pi$  mass spectrum includes a non-resonant scalar component.  
35 This value comes from a fit with  $\chi^2$  of 178/117.

### $K^*(892)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $K\pi$	~ 100	%
$\Gamma_2$ $(K\pi)^\pm$	( 99.901 ± 0.009 ) %	
$\Gamma_3$ $(K\pi)^0$	( 99.754 ± 0.021 ) %	
$\Gamma_4$ $K^0 \gamma$	( 2.46 ± 0.21 ) × 10 <sup>-3</sup>	
$\Gamma_5$ $K^\pm \gamma$	( 9.9 ± 0.9 ) × 10 <sup>-4</sup>	
$\Gamma_6$ $K\pi\pi$	< 7	× 10 <sup>-4</sup> 95%

See key on page 601

Meson Particle Listings

$K^*(892), K_1(1270)$

CONSTRAINED FIT INFORMATION		
An overall fit to the total width and a partial width uses 13 measurements and one constraint to determine 3 parameters. The overall fit has a $\chi^2 = 7.8$ for 11 degrees of freedom.		
The following <i>off-diagonal</i> array elements are the correlation coefficients $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters $p_i$ , including the branching fractions, $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the $x_i$ whose labels appear in this array to sum to one.		
$x_5$	$\begin{array}{ c } \hline -100 \\ \hline \end{array}$	
$\Gamma$	$\begin{array}{ c } \hline 19 \quad -19 \\ \hline \end{array}$	
	$x_2 \quad x_5$	
Mode	Rate (MeV)	
$\Gamma_2 \quad (K\pi)^\pm$	50.7 $\pm$ 0.9	
$\Gamma_5 \quad K^\pm \gamma$	0.050 $\pm$ 0.005	

CONSTRAINED FIT INFORMATION		
An overall fit to the total width and a partial width uses 22 measurements and one constraint to determine 3 parameters. The overall fit has a $\chi^2 = 66.8$ for 20 degrees of freedom.		
The following <i>off-diagonal</i> array elements are the correlation coefficients $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters $p_i$ , including the branching fractions, $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the $x_i$ whose labels appear in this array to sum to one.		
$x_4$	$\begin{array}{ c } \hline -100 \\ \hline \end{array}$	
$\Gamma$	$\begin{array}{ c } \hline 15 \quad -15 \\ \hline \end{array}$	
	$x_3 \quad x_4$	
Mode	Rate (MeV)	Scale factor
$\Gamma_3 \quad (K\pi)^0$	47.3 $\pm$ 0.6	2.1
$\Gamma_4 \quad K^0 \gamma$	0.116 $\pm$ 0.010	

$K^*(892)$ PARTIAL WIDTHS		
$\Gamma(K^0 \gamma)$		$\Gamma_4$
VALUE (keV)	EVTS	DOCUMENT ID
116 $\pm$ 10	OUR FIT	
116.5 $\pm$ 9.9	584	CARLSMITH 86 SPEC 0 $K_L^0 \rightarrow K_S^0 \pi^0 A$
$\Gamma(K^\pm \gamma)$		$\Gamma_5$
VALUE (keV)	DOCUMENT ID	TECN
50 $\pm$ 5	OUR FIT	
50 $\pm$ 5	OUR AVERAGE	
48 $\pm$ 11	BERG 83 SPEC -	156 $K^- \rightarrow \bar{K} \pi A$
51 $\pm$ 5	CHANDLEE 83 SPEC +	200 $K^+ \rightarrow K \pi A$

$K^*(892)$ BRANCHING RATIOS		
$\Gamma(K^0 \gamma) / \Gamma_{\text{total}}$		$\Gamma_4 / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN
2.46 $\pm$ 0.21	OUR FIT	
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1.5 $\pm$ 0.7	CARITHERS 75B CNTR 0	8-16 $\bar{K}^0 A$
$\Gamma(K^\pm \gamma) / \Gamma_{\text{total}}$		$\Gamma_5 / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN
0.99 $\pm$ 0.09	OUR FIT	
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<1.6	95	BEMPORAD 73 CNTR + 10-16 $K^+ A$
$\Gamma(K \pi \pi) / \Gamma((K \pi)^\pm)$		$\Gamma_6 / \Gamma_2$
VALUE	DOCUMENT ID	TECN
< 7 $\times 10^{-4}$	95	JONGEJANS 78 HBC 4 $K^- \rho \rightarrow \rho \bar{K}^0 2\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<20 $\times 10^{-4}$	WOJCICKI 64 HBC -	1.7 $K^- \rho \rightarrow \bar{K}^0 \pi^- \rho$

$K^*(892)$ REFERENCES		
LEES 13F	PR D87 052010	J.P. Lees <i>et al.</i> (BABAR Collab.)
DEL-AMO-SA... 11I	PR D83 072001	P. del Amo Sanchez <i>et al.</i> (BABAR Collab.)
ANTONELLI 10	EPJ C69 399	M. Antonelli <i>et al.</i> (FlaviaNet Working Group)
BOITO 10	JHEP 1009 031	D.R. Boito, R. Escribano, M. Jamin (BARC)
BOITO 09	EPJ C59 821	D.R. Boito, R. Escribano, M. Jamin
MITCHELL 09A	PR D79 072008	R.E. Mitchell <i>et al.</i> (CLEO Collab.)
BONVICINI 08A	PR D78 052001	G. Bonvicini <i>et al.</i> (CLEO Collab.)
JAMIN 08	PL B664 78	M. Jamin, A. Pich, J. Portoles
AUBERT 07AK	PR D76 012008	B. Aubert <i>et al.</i> (BABAR Collab.)
EPIFANOV 07	PL B654 65	D. Epifanov <i>et al.</i> (BELLE Collab.)
LINK 05I	PL B621 72	J.M. Link <i>et al.</i> (FNAL FOCUS Collab.)
BONVICINI 02	PRL 89 111803	G. Bonvicini <i>et al.</i> (CLEO Collab.)
PDG 00	EPJ C15 1	D.E. Groom <i>et al.</i> (PDG Collab.)
ABELE 99D	PL B468 178	A. Abele <i>et al.</i> (Crystal Barrel Collab.)
BARATE 99R	EPJ C11 599	R. Barate <i>et al.</i> (ALEPH Collab.)
BARBERIS 98E	PL B436 204	D. Barberis <i>et al.</i> (Omega Expt.)
BIRD 89	SLAC-332	P.F. Bird (SLAC)
ASTON 88	NP B296 493	D. Aston <i>et al.</i> (SLAC, NAGO, CINC, INUS)
ATKINSON 86	ZPHY C30 521	M. Atkinson <i>et al.</i> (BONN, CERN, GLAS+)
CARLSMITH 86	PRL 56 18	D. Carlsmith <i>et al.</i> (EFI, SACL)
BAUBILLIER 84B	ZPHY C26 37	M. Baubillier <i>et al.</i> (BIRM, CERN, GLAS+)
NAPIER 84	PL I49B 514	A. Napier <i>et al.</i> (TUFTS, ARIZ, FNAL, FLOR+)
BARTH 83	NP B223 296	M. Barth <i>et al.</i> (BRUX, CERN, GENO, MONS+)
BERG 83	Thesis UMI-83-21652	D.M. Berg (ROCH)
CHANDLEE 83	PRL 51 168	C. Chandlee <i>et al.</i> (ROCH, FNAL, MINN)
CLELAND 82	NP B208 189	W.E. Cleland <i>et al.</i> (DURH, GEVA, LAUS+)
DELFOSE 81	NP B183 349	A. Delfosse <i>et al.</i> (GEVA, LAUS)
TOAFF 81	PR D23 1500	S. Toaff <i>et al.</i> (ANL, KANS)
AJINENKO 80	ZPHY C5 177	I.V. Ajinenko <i>et al.</i> (SERP, BRUX, MONS+)
EVANGELIS... 80	NP B165 383	C. Evangelista <i>et al.</i> (BARI, BONN, CERN+)
AGUILAR... 78B	NP B141 101	M. Aguilar-Benitez <i>et al.</i> (MADR, TATA+)
BALAND 78	NP B140 220	J.F. Baland <i>et al.</i> (MONS, BELG, CERN+)
COOPER 78	NP B136 365	A.M. Cooper <i>et al.</i> (TATA, CERN, CDEF+)
JONGEJANS 78	NP B139 383	B. Jongejans <i>et al.</i> (ZEEM, CERN, NIJN+)
WICKLUND 78	PR D17 1197	A.B. Wicklund <i>et al.</i> (ANL)
BOWLER 77	NP B126 31	M.G. Bowler <i>et al.</i> (OXF)
CARITHERS 75B	PRL 35 349	W.C.J. Carithers <i>et al.</i> (ROCH, MCGI)
MCCUBBIN 75	NP B86 13	N.A. McCubbin, L. Lyons (OXF)
PALER 75	NP B96 1	K. Paler <i>et al.</i> (RHEL, SACL, EPOL)
FOX 74	NP B80 403	G.C. Fox, M.L. Griss (CIT)
MATISON 74	PR D9 1872	M.J. Matison <i>et al.</i> (LBL)
BEMPORAD 73	NP B51 1	C. Bemporad <i>et al.</i> (CERN, ETH, LOIC)
CLARK 73	NP B54 432	A.G. Clark, L. Lyons, D. Radojicic (OXF)
LEWIS 73	NP B60 283	P.H. Lewis <i>et al.</i> (LOWC, LOIC, CDEF)
LINGLIN 73	NP B55 408	D. Linglin (CERN)
BUCHNER 72	NP B45 333	K. Buchner <i>et al.</i> (MPIM, CERN, BRUX)
AGUILAR... 71B	PR D4 2583	M. Aguilar-Benitez, R.L. Eisner, J.B. Kinson (BNL)
HABER 70	NP B17 289	B. Haber <i>et al.</i> (REHO, SACL, BGNA, EPOL)
CRENNELL 69D	PRL 22 487	D.J. Crennell <i>et al.</i> (BNL)
DAVIS 69	PRL 23 1071	P.J. Davis <i>et al.</i> (LRL)
SCHWEING... 68	PR 166 1317	F. Schweingruber <i>et al.</i> (ANL, NWES)
BARASH 67B	PR 156 1399	N. Barash <i>et al.</i> (COLU)
BARLOW 67	NC 50A 701	J. Barlow <i>et al.</i> (CERN, CDEF, IRAD, LIVP)
DAUBER 67B	PR 153 1403	P.M. Dauber <i>et al.</i> (UCLA)
DEBAERE 67B	NC 51A 401	W. de Baere <i>et al.</i> (BRUX, CERN)
WOJCICKI 64	PR 135 B484	S.G. Wojcicki (LRL)

$K_1(1270)$

$I(J^P) = \frac{1}{2}(1^+)$

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### K<sub>1</sub>(1270) MASS

VALUE (MeV)	DOCUMENT ID		
<b>1272 ± 7 OUR AVERAGE</b>	Includes data from the 2 datablocks that follow this one.		

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### PRODUCED BY K<sup>-</sup>, BACKWARD SCATTERING, HYPERON EXCHANGE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1275 ± 10</b>	700	GAVILLET	78	HBC	+ 4.2 K <sup>-</sup> p → Ξ <sup>-</sup> (Kππ) <sup>+</sup>

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### PRODUCED BY K BEAMS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.				
<b>1270 ± 10</b>	<sup>1</sup> DAUM	81c	CNTR	- 63 K <sup>-</sup> p → K <sup>-</sup> 2πp
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 1276	<sup>2</sup> TORNQVIST	82b	RVUE	
~ 1300	VERGEEST	79	HBC	- 4.2 K <sup>-</sup> p → (K̄ππ) <sup>-</sup> p
1289 ± 25	<sup>3</sup> CARNEGIE	77	ASPK	± 13 K <sup>±</sup> p → (Kππ) <sup>±</sup> p
~ 1300	BRANDENB...	76	ASPK	± 13 K <sup>±</sup> p → (Kππ) <sup>±</sup> p
~ 1270	OTTER	76	HBC	- 10,14,16 K <sup>-</sup> p → (K̄ππ) <sup>-</sup> p
1260	DAVIS	72	HBC	+ 12 K <sup>+</sup> p
1234 ± 12	FIRESTONE	72b	DBC	+ 12 K <sup>+</sup> d

<sup>1</sup> Well described in the chiral unitary approach of GENG 07 with two poles at 1195 and 1284 MeV and widths of 246 and 146 MeV, respectively.

<sup>2</sup> From a unitarized quark-model calculation.

<sup>3</sup> From a model-dependent fit with Gaussian background to BRANDENBURG 76 data.

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### PRODUCED BY BEAMS OTHER THAN K MESONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1248.1 ± 3.3 ± 1.4</b>		GULER	11	BELL	B <sup>+</sup> → J/ψ K <sup>+</sup> π <sup>+</sup> π <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1279 ± 10	25k	<sup>4</sup> ABLIKIM	06c	BES2	J/ψ → K̄*(892) <sup>0</sup> K <sup>+</sup> π <sup>-</sup>
1294 ± 10	310	RODEBACK	81	HBC	4 π <sup>-</sup> p → Λ K 2π
1300	40	CRENNELL	72	HBC	0 4.5 π <sup>-</sup> p → Λ K 2π
1242 ± <sup>9</sup> / <sub>-10</sub>		<sup>5</sup> ASTIER	69	HBC	0 p̄p
1300	45	CRENNELL	67	HBC	0 6 π <sup>-</sup> p → Λ K 2π

<sup>4</sup> Systematic errors not estimated.

<sup>5</sup> This was called the C meson.

# Meson Particle Listings

## $K_1(1270)$

### PRODUCED IN $\tau$ LEPTON DECAYS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1254 \pm 33 \pm 34</math></b>	7k	ASNER	00B	CLEO	$\pm$ $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$

### $K_1(1270)$ WIDTH

VALUE (MeV)	DOCUMENT ID
<b><math>90 \pm 20</math> OUR ESTIMATE</b>	This is only an educated guess; the error given is larger than the error on the average of the published values.
<b><math>87 \pm 7</math> OUR AVERAGE</b>	Includes data from the 2 datablocks that follow this one.

### PRODUCED BY $K^-$ , BACKWARD SCATTERING, HYPERON EXCHANGE

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.					

<b><math>75 \pm 15</math></b>	700	GAVILLET	78	HBC	$+$ $4.2 K^- p \rightarrow \Xi^- K \pi \pi$
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### PRODUCED BY $K$ BEAMS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
The data in this block is included in the average printed for a previous datablock.				

<b><math>90 \pm 8</math></b>	<sup>6</sup> DAUM	81c	CNTR	$-$	$63 K^- p \rightarrow K^- 2\pi p$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$\sim 150$	VERGEEST	79	HBC	$-$	$4.2 K^- p \rightarrow (\overline{K} \pi \pi)^- p$
$150 \pm 71$	<sup>7</sup> CARNEGIE	77	ASPK	$\pm$	$13 K^\pm p \rightarrow (K \pi \pi)^\pm p$
$\sim 200$	BRANDENB...	76	ASPK	$\pm$	$13 K^\pm p \rightarrow (K \pi \pi)^\pm p$
120	DAVIS	72	HBC	$+$	$12 K^+ p$
$188 \pm 21$	FIRESTONE	72B	DBC	$+$	$12 K^+ d$
<sup>6</sup> Well described in the chiral unitary approach of GENG 07 with two poles at 1195 and 1284 MeV and widths of 246 and 146 MeV, respectively.					
<sup>7</sup> From a model-dependent fit with Gaussian background to BRANDENBURG 76 data.					

### PRODUCED BY BEAMS OTHER THAN $K$ MESONS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>119.5 \pm 5.2 \pm 6.7</math></b>		GULER	11	BELL	$B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$131 \pm 21$	25k	<sup>8</sup> ABLIKIM	06c	BES2	$J/\psi \rightarrow \overline{K}^*(892)^0 K^+ \pi^-$
$66 \pm 15$	310	RODEBACK	81	HBC	$4 \pi^- p \rightarrow \Lambda K 2\pi$
60	40	CRENNELL	72	HBC	0 $4.5 \pi^- p \rightarrow \Lambda K 2\pi$
$127 \pm_{-25}^7$		ASTIER	69	HBC	0 $\overline{p} p$
60	45	CRENNELL	67	HBC	0 $6 \pi^- p \rightarrow \Lambda K 2\pi$
<sup>8</sup> Systematic errors not estimated.					

### PRODUCED IN $\tau$ LEPTON DECAYS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>260 \pm_{-70}^{90} \pm 80</math></b>	7k	ASNER	00B	CLEO	$\pm$ $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$

### $K_1(1270)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 K \rho$	( $42 \pm 6$ ) %
$\Gamma_2 K_0^*(1430) \pi$	( $28 \pm 4$ ) %
$\Gamma_3 K^*(892) \pi$	( $16 \pm 5$ ) %
$\Gamma_4 K \omega$	( $11.0 \pm 2.0$ ) %
$\Gamma_5 K \rho_0(1370)$	( $3.0 \pm 2.0$ ) %
$\Gamma_6 \gamma K^0$	seen

### $K_1(1270)$ PARTIAL WIDTHS

$\Gamma(K\rho)$					$\Gamma_1$
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
57 ± 5	MAZZUCATO	79	HBC	+ 4.2 $K^- p \rightarrow \Xi^- (K\pi\pi)^+$	
75 ± 6	CARNEGIE	77B	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$	
$\Gamma(K_0^*(1430)\pi)$					$\Gamma_2$
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
26 ± 6	CARNEGIE	77B	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$	
$\Gamma(K^*(892)\pi)$					$\Gamma_3$
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
14 ± 11	MAZZUCATO	79	HBC	+ 4.2 $K^- p \rightarrow \Xi^- (K\pi\pi)^+$	
2 ± 2	CARNEGIE	77B	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$	
$\Gamma(K\omega)$					$\Gamma_4$
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4 ± 4	MAZZUCATO	79	HBC	+ 4.2 $K^- p \rightarrow \Xi^- (K\pi\pi)^+$	
24 ± 3	CARNEGIE	77B	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$	

### $\Gamma(K \rho_0(1370))$

$\Gamma_5$

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$22 \pm 5$	CARNEGIE	77B	ASPK	$\pm$ $13 K^\pm p \rightarrow (K \pi \pi)^\pm p$

### $\Gamma(\gamma K^0)$

$\Gamma_6$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b><math>73.2 \pm 6.1 \pm 28.3</math></b>	ALAVI-HARATI	02B	KTEV $K^+ A \rightarrow K^* + A$

### $K_1(1270)$ BRANCHING RATIOS

$\Gamma(K\rho)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.42 ± 0.06</b>	<sup>9</sup> DAUM	81c	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.584 ± 0.043	<sup>10</sup> GULER	11	BELL $B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$
dominant	RODEBACK	81	HBC $4 \pi^- p \rightarrow \Lambda K 2\pi$

### $\Gamma(K_0^*(1430) \pi)/\Gamma_{\text{total}}$

$\Gamma_2/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.28 \pm 0.04</math></b>	<sup>9</sup> DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
$0.0201 \pm 0.0064$	<sup>10</sup> GULER	11	BELL $B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$

### $\Gamma(K^*(892) \pi)/\Gamma_{\text{total}}$

$\Gamma_3/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.16 \pm 0.05</math></b>	<sup>9</sup> DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
$0.171 \pm 0.023$	<sup>10</sup> GULER	11	BELL $B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$

### $\Gamma(K \omega)/\Gamma_{\text{total}}$

$\Gamma_4/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.11 \pm 0.02</math></b>	<sup>9</sup> DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
$0.225 \pm 0.052$	<sup>10</sup> GULER	11	BELL $B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$

### $\Gamma(K \omega)/\Gamma(K \rho)$

$\Gamma_4/\Gamma_1$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$< 0.30$	95	RODEBACK	81	HBC $4 \pi^- p \rightarrow \Lambda K 2\pi$

### $\Gamma(K \rho_0(1370))/\Gamma_{\text{total}}$

$\Gamma_5/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.03 \pm 0.02</math></b>	<sup>9</sup> DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$

### D-wave/S-wave RATIO FOR $K_1(1270) \rightarrow K^*(892) \pi$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.0 \pm 0.7</math></b>	<sup>9</sup> DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$
<sup>9</sup> Average from low and high $t$ data.			
<sup>10</sup> Assuming that decays are saturated by the $K \rho$ , $K_0^*(1430) \pi$ , $K^*(892) \pi$ , $K \omega$ decay modes and neglecting interference between them. The values $B(\omega \rightarrow \pi^+ \pi^-) = (1.53 \pm_{-0.13}^{+0.11})\%$ and $B(K_0^*(1430) \rightarrow K \pi) = (93 \pm 10)\%$ are used. Systematic uncertainties not estimated.			

### $K_1(1270)$ REFERENCES

GULER	11	PR D83 032005	H. Guler <i>et al.</i>	(BELLE Collab.)
GENG	07	PR D75 014017	L.S. Geng <i>et al.</i>	
ABLIKIM	06C	PL B633 681	M. Ablikim <i>et al.</i>	(BES Collab.)
ALAVI-HARATI	02B	PRL 89 072001	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
ASNER	00B	PR D62 072006	D.M. Asner <i>et al.</i>	(CLEO Collab.)
TORNGVIST	82B	NP B203 268	N.A. Tornqvist	(HEL5)
DAUM	81C	NP B187 1	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
RODEBACK	81	ZPHY C9 9	S. Rodeback <i>et al.</i>	(CERN, CDEF, MADR+)
MAZZUCATO	79	NP B156 532	M. Mazzucato <i>et al.</i>	(CERN, ZEEM, NUM+)
VERGEEST	79	NP B158 265	J.S.M. Vergeest <i>et al.</i>	(NUM, AMST, CERN+)
GAVILLET	78	PL 76B 517	P. Gavillet <i>et al.</i>	(AMST, CERN, NUM+)
CARNEGIE	77	NP B127 509	R.K. Carnegie <i>et al.</i>	(SLAC)
CARNEGIE	77B	PL 68B 287	R.K. Carnegie <i>et al.</i>	(SLAC)
BRANDENB...	76	PRL 36 703	G.W. Brandenburg <i>et al.</i>	(SLAC)JP
OTTER	76	NP B106 77	G. Otter <i>et al.</i>	(AACH3, BERL, CERN, LOIC+)
CRENNELL	72	PR D6 1220	D.J. Crennell <i>et al.</i>	(BNL)
DAVIS	72	PR D5 2688	P.J. Davis <i>et al.</i>	(LBL)
FIRESTONE	72B	PR D5 505	A. Firestone <i>et al.</i>	(LBL)
ASTIER	69	NP B10 65	A. Astier <i>et al.</i>	(CDEF, CERN, IPNP, LIPV)IJP
CRENNELL	67	PRL 19 44	D.J. Crennell <i>et al.</i>	(BNL)I



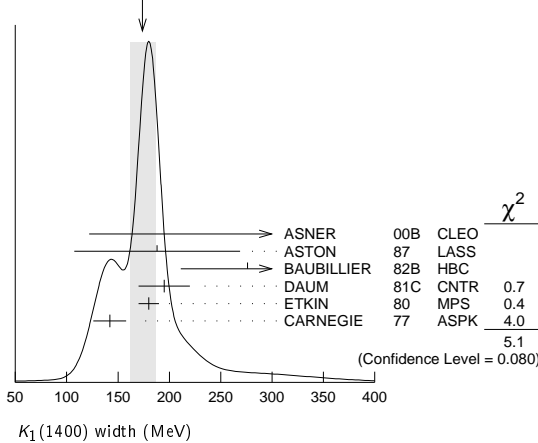
See key on page 601

## Meson Particle Listings

 $K_1(1400)$ ,  $K^*(1410)$ 

$K_1(1400)$ MASS					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b>1403 ± 7 OUR AVERAGE</b>					
1463 ± 64 ± 68	7k	ASNER	00B	CLEO	± $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$
1373 ± 14 ± 18		<sup>1</sup> ASTON	87	LASS	0 11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
1392 ± 18		BAUBILLIER	82B	HBC	0 8.25 $K^- p \rightarrow K_S^0 \pi^+ \pi^- n$
1410 ± 25		DAUM	81C	CNTR	— 63 $K^- p \rightarrow K^- 2\pi p$
1415 ± 15		ETKIN	80	MPS	0 6 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
1404 ± 10		<sup>2</sup> CARNEGIE	77	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1418 ± 8	25k	<sup>3</sup> ABLIKIM	06C	BES2	$J/\psi \rightarrow \bar{K}^*(892)^0 K^\pm \pi^-$
~ 1350		<sup>4</sup> TORNQVIST	82B	RVUE	
~ 1400		VERGEEST	79	HBC	— 4.2 $K^- p \rightarrow (\bar{K}\pi\pi)^- p$
~ 1400		BRANDENB...	76	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
1420		DAVIS	72	HBC	+ 12 $K^+ p$
1368 ± 18		FIRESTONE	72B	DBC	+ 12 $K^+ d$
<sup>1</sup> From partial-wave analysis of $K^0 \pi^+ \pi^-$ system. <sup>2</sup> From a model-dependent fit with Gaussian background to BRANDENBURG 76 data. <sup>3</sup> Systematic errors not estimated. <sup>4</sup> From a unitarized quark-model calculation.					

$K_1(1400)$ WIDTH					
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
<b>174 ± 13 OUR AVERAGE</b> Error includes scale factor of 1.6. See the ideogram below.					
300 + 370 - 110 ± 140	7k	ASNER	00B	CLEO	± $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$
188 ± 54 ± 60		<sup>5</sup> ASTON	87	LASS	0 11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
276 ± 65		BAUBILLIER	82B	HBC	0 8.25 $K^- p \rightarrow K_S^0 \pi^+ \pi^- n$
195 ± 25		DAUM	81C	CNTR	— 63 $K^- p \rightarrow K^- 2\pi p$
180 ± 10		ETKIN	80	MPS	0 6 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
142 ± 16		<sup>6</sup> CARNEGIE	77	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
152 ± 16	25k	<sup>7</sup> ABLIKIM	06C	BES2	$J/\psi \rightarrow \bar{K}^*(892)^0 K^\pm \pi^-$
~ 200		VERGEEST	79	HBC	— 4.2 $K^- p \rightarrow (\bar{K}\pi\pi)^- p$
~ 160		BRANDENB...	76	ASPK	± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
80		DAVIS	72	HBC	+ 12 $K^+ p$
241 ± 30		FIRESTONE	72B	DBC	+ 12 $K^+ d$
<sup>5</sup> From partial-wave analysis of $K^0 \pi^+ \pi^-$ system. <sup>6</sup> From a model-dependent fit with Gaussian background to BRANDENBURG 76 data. <sup>7</sup> Systematic errors not estimated.					

WEIGHTED AVERAGE  
174±13 (Error scaled by 1.6)

$K_1(1400)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $K^*(892)\pi$	(94 ± 6) %	
$\Gamma_2$ $K\rho$	(3.0 ± 3.0) %	
$\Gamma_3$ $K f_0(1370)$	(2.0 ± 2.0) %	
$\Gamma_4$ $K\omega$	(1.0 ± 1.0) %	
$\Gamma_5$ $K_S^0(1430)\pi$	not seen	
$\Gamma_6$ $\gamma K^0$	seen	

$K_1(1400)$ PARTIAL WIDTHS				
$\Gamma(K^*(892)\pi)$	VALUE (MeV)	DOCUMENT ID	TECN	CHG COMMENT
<b>117 ± 10</b>		CARNEGIE	77	ASPK ± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
$\Gamma(K\rho)$	VALUE (MeV)	DOCUMENT ID	TECN	CHG COMMENT
<b>2 ± 1</b>		CARNEGIE	77	ASPK ± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
$\Gamma(K\omega)$	VALUE (MeV)	DOCUMENT ID	TECN	CHG COMMENT
<b>23 ± 12</b>		CARNEGIE	77	ASPK ± 13 $K^\pm p \rightarrow (K\pi\pi)^\pm p$
$\Gamma(\gamma K^0)$	VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>280.8 ± 23.2 ± 40.4</b>		ALAVI-HARATI02B	KTEV	$K + A \rightarrow K^* + A$

$K_1(1400)$ BRANCHING RATIOS				
$\Gamma(K^*(892)\pi)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.94 ± 0.06</b>		<sup>8</sup> DAUM	81C	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
$\Gamma(K\rho)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.03 ± 0.03</b>		<sup>8</sup> DAUM	81C	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
$\Gamma(K f_0(1370))/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.02 ± 0.02</b>		<sup>8</sup> DAUM	81C	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
$\Gamma(K\omega)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.01 ± 0.01</b>		<sup>8</sup> DAUM	81C	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
$\Gamma(K_S^0(1430)\pi)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT
not seen		<sup>8</sup> DAUM	81C	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
D-wave/S-wave RATIO FOR $K_1(1400) \rightarrow K^*(892)\pi$				
$D/S$	VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.04 ± 0.01</b>		<sup>8</sup> DAUM	81C	CNTR 63 $K^- p \rightarrow K^- 2\pi p$
<sup>8</sup> Average from low and high $t$ data.				

$K_1(1400)$ REFERENCES				
ABLIKIM	06C	PL B633 681	M. Ablikim <i>et al.</i>	(BES Collab.)
ALAVI-HARATI	02B	PRL 89 072001	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
ASNER	00B	PR D62 072006	D.M. Asner <i>et al.</i>	(CLEO Collab.)
ASTON	87	NP B292 693	D. Aston <i>et al.</i>	(SLAC, NAGO, CIN, INUS)
BAUBILLIER	82B	NP B202 21	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
TORNQVIST	82B	NP B203 268	N.A. Tornqvist	(HELS)
DAUM	81C	NP B187 1	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
ETKIN	80	PR D22 42	A. Etkin <i>et al.</i>	(BNL, CUNY) JP
VERGEEST	79	NP B158 265	J.S.M. Vergeest <i>et al.</i>	(NUM, AMST, CERN+)
CARNEGIE	77	NP B127 509	R.K. Carnegie <i>et al.</i>	(SLAC)
BRANDENB...	76	PRL 36 703	G.W. Brandenburg <i>et al.</i>	(SLAC) JP
DAVIS	72	PR D5 2688	P.J. Davis <i>et al.</i>	(LBL)
FIRESTONE	72B	PR D5 505	A. Firestone <i>et al.</i>	(LBL)

$K^*(1410)$ MASS					
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>1414 ± 15 OUR AVERAGE</b> Error includes scale factor of 1.3.					
1380 ± 21 ± 19	ASTON	88	LASS	0	11 $K^- p \rightarrow K^- \pi^+ n$
1420 ± 7 ± 10	ASTON	87	LASS	0	11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1276 + 72 - 77	<sup>1,2</sup> BOITO	09	RVUE		$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
1367 ± 54	BIRD	89	LASS	—	11 $K^- p \rightarrow \bar{K}^0 \pi^- p$
1474 ± 25	BAUBILLIER	82B	HBC	0	8.25 $K^- p \rightarrow \bar{K}^0 2\pi n$
1500 ± 30	ETKIN	80	MPS	0	6 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
<sup>1</sup> From the pole position of the $K\pi$ vector form factor in the complex $s$ -plane and using EPIFANOV 07 data. <sup>2</sup> Systematic uncertainties not estimated.					

$K^*(1410)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>232 ± 21 OUR AVERAGE</b> Error includes scale factor of 1.1.				
176 ± 52 ± 22	ASTON	88	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$
240 ± 18 ± 12	ASTON	87	LASS	0 11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$

# Meson Particle Listings

## $K^*(1410)$ , $K_0^*(1430)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$198^{+61}_{-87}$	3.4 BOITO	09	RVUE	$\tau^- \rightarrow K_S^0 \pi^- \nu_\tau$
$114 \pm 101$	BIRD	89	LASS	$-$ 11 $K^- p \rightarrow \overline{K}^0 \pi^- p$
$275 \pm 65$	BAUBILLIER	82B	HBC	0 8.25 $K^- p \rightarrow \overline{K}^0 2\pi n$
$500 \pm 100$	ETKIN	80	MPS	0 6 $K^- p \rightarrow \overline{K}^0 \pi^+ \pi^- n$

<sup>3</sup> From the pole position of the  $K\pi$  vector form factor in the complex  $s$ -plane and using EPIFANOV 07 data.  
<sup>4</sup> Systematic uncertainties not estimated.

$K^*(1410)$ DECAY MODES				
Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level		
$\Gamma_1$ $K^*(892)\pi$	$> 40$	%		
$\Gamma_2$ $K\pi$	( $6.6 \pm 1.3$ )	%		
$\Gamma_3$ $K\rho$	$< 7$	%		
$\Gamma_4$ $\gamma K^0$	seen			

$K^*(1410)$ PARTIAL WIDTHS				
$\Gamma(\gamma K^0)$				$\Gamma_4$
VALUE (keV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;52.9</b>	90	ALAVI-HARATI02B	KTEV	$K + A \rightarrow K^* + A$

K*(1410) BRANCHING RATIOS						
$\Gamma(K\rho)/\Gamma(K^*(892)\pi)$		$\Gamma_3/\Gamma_1$				
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.17	95	ASTON	84	LASS	0	11 $K^- p \rightarrow \overline{K}^0 2\pi n$

$\Gamma(K\pi)/\Gamma(K^*(892)\pi)$					$\Gamma_2/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<0.16	95	ASTON	84	LASS	0 11 $K^- p \rightarrow \overline{K}^0 2\pi n$

$\Gamma(K\pi)/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.066 \pm 0.010 \pm 0.008</math></b>		ASTON	88	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$

$K^*(1410)$ REFERENCES				
BOITO	09	EPI C59 821	D.R. Boito, R. Escribano, M. Jamin	
EPIFANOV	07	PL B654 65	D. Epifanov <i>et al.</i>	(BELLE Collab.)
ALAVI-HARATI	02B	PRL 89 072001	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
BIRD	89	SLAC-332	P.F. Bird	(SLAC)
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	87	NP B292 693	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	84	PL 149B 258	D. Aston <i>et al.</i>	(SLAC, CARL, OTTA)JP
BAUBILLIER	82B	NP B202 21	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
ETKIN	80	PR D22 42	A. Etkin <i>et al.</i>	(BNL, CUNY)JP

<div><math>K_0^*(1430)</math></div>	$I(J^P) = \frac{1}{2}(0^+)$
See our minireview in the 1994 edition and in this edition under the $f_0(500)$ .	

$K_0^*(1430)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1425 <math>\pm</math> 50 OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1438 $\pm$ 8 $\pm$ 4	5.4k	<sup>1</sup> LEES	14E	BABR $\eta_C(1S) \rightarrow K^+ K^- \eta/\pi^0$
1427 $\pm$ 4 $\pm$ 13		<sup>2</sup> BUGG	10	RVUE S-matrix pole
1466.6 $\pm$ 0.7 $\pm$ 3.4	141k	<sup>3</sup> BONVICINI	08A	CLEO $D^+ \rightarrow K^- \pi^+ \pi^+$
$\sim$ 1412		<sup>4</sup> LINK	07	FOCS $D^+ \rightarrow K^- K^+ \pi^+$
1461.0 $\pm$ 4.0 $\pm$ 2.1	54k	<sup>5</sup> LINK	07B	FOCS $D^+ \rightarrow K^- \pi^+ \pi^+$
1406 $\pm$ 29		<sup>6</sup> BUGG	06	RVUE
1435 $\pm$ 6		<sup>7</sup> ZHOU	06	RVUE $K\rho \rightarrow K^- \pi^+ n$
1455 $\pm$ 20 $\pm$ 15		ABLIKIM	05Q	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^-$
1456 $\pm$ 8		<sup>8</sup> ZHENG	04	RVUE $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 1419		<sup>9</sup> BUGG	03	RVUE 11 $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 1440		<sup>10</sup> LI	03	RVUE 11 $K^- p \rightarrow K^- \pi^+ n$
1459 $\pm$ 9	15k	<sup>11</sup> AITALA	02	E791 $D^+ \rightarrow K^- \pi^+ \pi^+$
$\sim$ 1440		<sup>12</sup> JAMIN	00	RVUE $K\rho \rightarrow K\rho$
1436 $\pm$ 8		<sup>13</sup> BARBERIS	98E	OMEG 450 $p\rho \rightarrow p_f p_s K^+ K^- \pi^+ \pi^-$
1415 $\pm$ 25		<sup>9</sup> ANISOVICH	97C	RVUE 11 $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 1450		<sup>14</sup> TORNQVIST	96	RVUE $\pi\pi \rightarrow \pi\pi, K\overline{K}, K\pi$
1412 $\pm$ 6		<sup>15</sup> ASTON	88	LASS 11 $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 1430		BAUBILLIER	84B	HBC 8.25 $K^- p \rightarrow \overline{K}^0 \pi^- p$
$\sim$ 1425		<sup>16</sup> ESTABROOKS	78	ASPK 13 $K^\pm p \rightarrow K^\pm \pi^\pm(n, \Delta)$
$\sim$ 1450.0		MARTIN	78	SPEC 10 $K^\pm p \rightarrow K_S^0 \pi p$

- Using both  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+ \pi^- \pi^0$ . From a likelihood scan in the presence of several interfering scalar-meson resonances with fixed width  $\Gamma(K_0^*(1430)) = 210$  MeV.
- S-Matrix pole. Supersedes BUGG 06. Combined analysis of ASTON 88, ABLIKIM 06c, AITALA 06, and LINK 09 using an  $s$ -dependent width with couplings to  $K\pi$  and  $K\eta'$ , and the Adler zero near thresholds.
- From the isobar model with a complex pole for the  $\kappa$ .
- From a non-parametric analysis.
- A Breit-Wigner mass and width.
- S-matrix pole. Reanalysis of ASTON 88, AITALA 02, and ABLIKIM 06c including the  $\kappa$  with an  $s$ -dependent width and an Adler zero near threshold.
- S-matrix pole. Using ASTON 88 and assuming  $K_0^*(800)$ ,  $K_0^*(1950)$ .
- Using ASTON 88 and assuming  $K_0^*(800)$ .
- T-matrix pole. Reanalysis of ASTON 88 data.
- Breit-Wigner fit. Using ASTON 88.
- Assuming a low-mass scalar  $K\pi$  resonance,  $\kappa(800)$ .
- T-matrix pole. Using data from ESTABROOKS 78 and ASTON 88.
- $J^P$  not determined, could be  $K_2^*(1430)$ .
- T-matrix pole.
- Uses a model for the background, without this background they get a mass 1340 MeV, where the phase shift passes 90°.
- Mass defined by pole position. From elastic  $K\pi$  partial-wave analysis.

$K_0^*(1430)$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>270 <math>\pm</math> 80 OUR ESTIMATE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
210 $\pm$ 20 $\pm$ 12	5.4k	<sup>1</sup> LEES	14E	BABR $\eta_C(1S) \rightarrow K^+ K^- \eta/\pi^0$
270 $\pm$ 10 $\pm$ 40		<sup>2</sup> BUGG	10	RVUE S-matrix pole
174.2 $\pm$ 1.9 $\pm$ 3.2	141k	<sup>3</sup> BONVICINI	08A	CLEO $D^+ \rightarrow K^- \pi^+ \pi^+$
$\sim$ 500		<sup>4</sup> LINK	07	FOCS $D^+ \rightarrow K^- K^+ \pi^+$
177.0 $\pm$ 8.0 $\pm$ 3.4	54k	<sup>5</sup> LINK	07B	FOCS $D^+ \rightarrow K^- \pi^+ \pi^+$
350 $\pm$ 40		<sup>6</sup> BUGG	06	RVUE
288 $\pm$ 22		<sup>7</sup> ZHOU	06	RVUE $K\rho \rightarrow K^- \pi^+ n$
270 $\pm$ 45 $\pm$ 30 $\pm$ 35		ABLIKIM	05Q	BES2 $\psi(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^-$
217 $\pm$ 31		<sup>8</sup> ZHENG	04	RVUE $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 316		<sup>9</sup> BUGG	03	RVUE 11 $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 350		<sup>10</sup> LI	03	RVUE 11 $K^- p \rightarrow K^- \pi^+ n$
175 $\pm$ 17	15k	<sup>11</sup> AITALA	02	E791 $D^+ \rightarrow K^- \pi^+ \pi^+$
$\sim$ 300		<sup>12</sup> JAMIN	00	RVUE $K\rho \rightarrow K\rho$
196 $\pm$ 45		<sup>13</sup> BARBERIS	98E	OMEG 450 $p\rho \rightarrow p_f p_s K^+ K^- \pi^+ \pi^-$
330 $\pm$ 50		<sup>9</sup> ANISOVICH	97C	RVUE 11 $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 320		<sup>14</sup> TORNQVIST	96	RVUE $\pi\pi \rightarrow \pi\pi, K\overline{K}, K\pi$
294 $\pm$ 23		ASTON	88	LASS 11 $K^- p \rightarrow K^- \pi^+ n$
$\sim$ 200		BAUBILLIER	84B	HBC 8.25 $K^- p \rightarrow \overline{K}^0 \pi^- p$
200 to 300		<sup>15</sup> ESTABROOKS	78	ASPK 13 $K^\pm p \rightarrow K^\pm \pi^\pm(n, \Delta)$

- Using both  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+ \pi^- \pi^0$ . From a likelihood scan in the presence of several interfering scalar-meson resonances with fixed mass  $M(K_0^*(1430)) = 1435$  MeV.
- S-Matrix pole. Supersedes BUGG 06. Combined analysis of ASTON 88, ABLIKIM 06c, AITALA 06, and LINK 09 using an  $s$ -dependent width with couplings to  $K\pi$  and  $K\eta'$ , and the Adler zero near thresholds.
- From the isobar model with a complex pole for the  $\kappa$ .
- From a non-parametric analysis.
- A Breit-Wigner mass and width.
- S-matrix pole. Reanalysis of ASTON 88, AITALA 02, and ABLIKIM 06c including the  $\kappa$  with an  $s$ -dependent width and an Adler zero near threshold.
- S-matrix pole. Using ASTON 88 and assuming  $K_0^*(800)$ ,  $K_0^*(1950)$ .
- Using ASTON 88 and assuming  $K_0^*(800)$ .
- T-matrix pole. Reanalysis of ASTON 88 data.
- Breit-Wigner fit. Using ASTON 88.
- Assuming a low-mass scalar  $K\pi$  resonance,  $\kappa(800)$ .
- T-matrix pole. Using data from ESTABROOKS 78 and ASTON 88.
- $J^P$  not determined, could be  $K_2^*(1430)$ .
- T-matrix pole.
- From elastic  $K\pi$  partial-wave analysis.

$K_0^*(1430)$ DECAY MODES				
Mode	Fraction ( $\Gamma_i/\Gamma$ )			
$\Gamma_1$ $K\pi$	(93 $\pm$ 10 )	%		
$\Gamma_2$ $K\eta$	( 8.6 $\pm$ 2.7 $\pm$ 3.4 )	%		
$\Gamma_3$ $K\eta'(958)$	seen			

$K_0^*(1430)$ BRANCHING RATIOS				
$\Gamma(K\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.93 \pm 0.04 \pm 0.09</math></b>	ASTON	88	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$

See key on page 601

## Meson Particle Listings

 $K_0^*(1430), K_2^*(1430)$  $\Gamma(K\eta)/\Gamma(K\pi)$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
$9.2 \pm 2.5^{+1.0}_{-2.5}$	5.4k	<sup>1</sup> LEES	14E BABR	$\eta_c(1S) \rightarrow K^+ K^- \eta/\pi^0$

<sup>1</sup> Using both  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$ . From a Dalitz analysis in the presence of several interfering scalar-meson resonances.

 $\Gamma(K\eta'(958))/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
seen	ABLIKIM	14J BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$

 $\Gamma_2/\Gamma_1$  $\Gamma_3/\Gamma$  $K_2^*(1430)$  REFERENCES

ABLIKIM	14J	PR D89 074030	M. Ablikim <i>et al.</i>	(BES III Collab.)
LEES	14E	PR D89 112004	J.P. Lees <i>et al.</i>	(BABAR Collab.)
BUGG	10	PR D81 014002	D.V. Bugg	(LOQM)
LINK	09	PL B681 14	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
BONVICINI	08A	PR D78 052001	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
LINK	07	PL B648 156	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	07B	PL B653 1	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ABLIKIM	06C	PL B633 681	M. Ablikim <i>et al.</i>	(BES Collab.)
AITALA	06	PR D73 032004	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
Also		PR D74 059901 (errat.)	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BUGG	06	PL B632 471	D.V. Bugg	(LOQM)
ZHOU	06	NP A775 212	Z.Y. Zhou, H.Q. Zheng	
ABLIKIM	05Q	PR D72 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
ZHENG	04	NP A733 235	H.Q. Zheng <i>et al.</i>	
BUGG	03	PL B572 1	D.V. Bugg	
LI	03	PR D67 034025	L. Li, B. Zou, G. Li	
AITALA	02	PRL 89 121801	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
JAMIN	00	NP B587 331	M. Jamin <i>et al.</i>	(Omega Expt.)
BARBERIS	98E	PL B436 204	D. Barberis <i>et al.</i>	
ANISOVICH	97C	PL B413 137	A.V. Anisovich, A.V. Sarantsev	
TORNQVIST	96	PRL 76 1575	N.A. Tornqvist, M. Roos	(HELs)
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
BAUBILLIER	84B	ZPHY C26 37	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
ESTABROOKS	78	NP B133 490	P.G. Estabrooks <i>et al.</i>	(MCGI, CARL, DURH+)
MARTIN	78	NP B134 392	A.D. Martin <i>et al.</i>	(DURH, GEVA)

 $K_2^*(1430)$ 

$$J(P) = \frac{1}{2}(2^+)$$

We consider that phase-shift analyses provide more reliable determinations of the mass and width.

 $K_2^*(1430)$  MASSCHARGED ONLY, WITH FINAL STATE  $K\pi$ 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$1425.6 \pm 1.5$	OUR AVERAGE	Error includes scale factor of 1.1.			
1420 $\pm 4$	1587	BAUBILLIER	84B HBC	—	$8.25 K^- p \rightarrow \bar{K}^0 \pi^- p$
1436 $\pm 5.5$	400	<sup>1,2</sup> CLELAND	82 SPEC	+	$30 K^+ p \rightarrow K_S^0 \pi^+ p$
1430 $\pm 3.2$	1500	<sup>1,2</sup> CLELAND	82 SPEC	+	$50 K^+ p \rightarrow K_S^0 \pi^+ p$
1430 $\pm 3.2$	1200	<sup>1,2</sup> CLELAND	82 SPEC	—	$50 K^+ p \rightarrow K_S^0 \pi^- p$
1423 $\pm 5$	935	TOAFF	81 HBC	—	$6.5 K^- p \rightarrow \bar{K}^0 \pi^- p$
1428.0 $\pm 4.6$		<sup>3</sup> MARTIN	78 SPEC	+	$10 K^\pm p \rightarrow K_S^0 \pi p$
1423.8 $\pm 4.6$		<sup>3</sup> MARTIN	78 SPEC	—	$10 K^\pm p \rightarrow K_S^0 \pi p$
1420.0 $\pm 3.1$	1400	AGUILAR-...	71B HBC	—	$3.9, 4.6 K^- p$
1425 $\pm 8.0$	225	<sup>1,2</sup> BARNHAM	71C HBC	+	$K^+ p \rightarrow K^0 \pi^+ p$
1416 $\pm 10$	220	CRENNELL	69D DBC	—	$3.9 K^- N \rightarrow \bar{K}^0 \pi^- N$
1414 $\pm 13.0$	60	<sup>1</sup> LIND	69 HBC	+	$9 K^+ p \rightarrow K^0 \pi^+ p$
1427 $\pm 12$	63	<sup>1</sup> SCHWEING...	68 HBC	—	$5.5 K^- p \rightarrow \bar{K}^0 \pi^- N$
1423 $\pm 11.0$	39	<sup>1</sup> BASSANO	67 HBC	—	$4.6-5.0 K^- p \rightarrow \bar{K}^0 \pi^- p$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1423.4 $\pm 2 \pm 3$	24809 $\pm 820$	<sup>4</sup> BIRD	89 LASS	—	$11 K^- p \rightarrow \bar{K}^0 \pi^- p$
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## NEUTRAL ONLY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$1432.4 \pm 1.3$	OUR AVERAGE			
1431.2 $\pm 1.8 \pm 0.7$		<sup>5</sup> ASTON	88 LASS	$11 K^- p \rightarrow K^- \pi^+ n$
1434 $\pm 4 \pm 6$		<sup>5</sup> ASTON	87 LASS	$11 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
1433 $\pm 6 \pm 10$		<sup>5</sup> ASTON	84B LASS	$11 K^- p \rightarrow \bar{K}^0 2\pi n$
1471 $\pm 12$		<sup>5</sup> BAUBILLIER	82B HBC	$8.25 K^- p \rightarrow N K_S^0 \pi \pi$
1428 $\pm 3$		<sup>5</sup> ASTON	81C LASS	$11 K^- p \rightarrow K^- \pi^+ n$
1434 $\pm 2$		<sup>5</sup> ESTABROOKS	78 ASPK	$13 K^\pm p \rightarrow p K \pi$
1440 $\pm 10$		<sup>5</sup> BOWLER	77 DBC	$5.5 K^+ d \rightarrow K \pi p p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1428.5 $\pm 3.9$	1786 $\pm 127$	<sup>6</sup> AUBERT	07AK BABR	$10.6 e^+ e^- \rightarrow K^{*0} K^\pm \pi^\mp \gamma$
1420 $\pm 7$	300	HENDRICK	76 DBC	$8.25 K^+ N \rightarrow K^+ \pi N$
1421.6 $\pm 4.2$	800	MCCUBBIN	75 HBC	$3.6 K^- p \rightarrow K^- \pi^+ n$
1420.1 $\pm 4.3$		<sup>7</sup> LINGLIN	73 HBC	$2-13 K^+ p \rightarrow K^+ \pi^- X$
1419.1 $\pm 3.7$	1800	AGUILAR-...	71B HBC	$3.9, 4.6 K^- p$
1416 $\pm 6$	600	CORDS	71 DBC	$9 K^+ n \rightarrow K^+ \pi^- p$
1421.1 $\pm 2.6$	2200	DAVIS	69 HBC	$12 K^+ p \rightarrow K^+ \pi^- X$

<sup>1</sup> Errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>2</sup> Number of events in peak re-evaluated by us.

<sup>3</sup> Systematic error added by us.

<sup>4</sup> From a partial wave amplitude analysis.

<sup>5</sup> From phase shift or partial-wave analysis.

<sup>6</sup> Systematic errors not estimated.

<sup>7</sup> From pole extrapolation, using world  $K^+ p$  data summary tape.

 $K_2^*(1430)$  WIDTHCHARGED ONLY, WITH FINAL STATE  $K\pi$ 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$98.5 \pm 2.7$	OUR FIT	Error includes scale factor of 1.1.			
$98.5 \pm 2.9$	OUR AVERAGE	Error includes scale factor of 1.1.			
109 $\pm 22$	400	<sup>8,9</sup> CLELAND	82 SPEC	+	$30 K^+ p \rightarrow K_S^0 \pi^+ p$
124 $\pm 12.8$	1500	<sup>8,9</sup> CLELAND	82 SPEC	+	$50 K^+ p \rightarrow K_S^0 \pi^+ p$
113 $\pm 12.8$	1200	<sup>8,9</sup> CLELAND	82 SPEC	—	$50 K^+ p \rightarrow K_S^0 \pi^- p$
85 $\pm 16$	935	TOAFF	81 HBC	—	$6.5 K^- p \rightarrow \bar{K}^0 \pi^- p$
96.5 $\pm 3.8$		MARTIN	78 SPEC	+	$10 K^\pm p \rightarrow K_S^0 \pi p$
97.7 $\pm 4.0$		MARTIN	78 SPEC	—	$10 K^\pm p \rightarrow K_S^0 \pi p$
94.7 $\pm 15.1$ —12.5	1400	AGUILAR-...	71B HBC	—	$3.9, 4.6 K^- p$

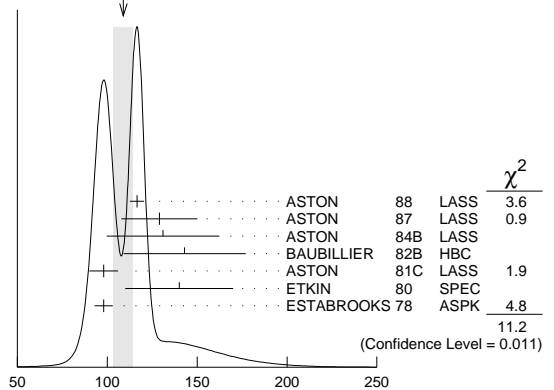
• • • We do not use the following data for averages, fits, limits, etc. • • •

98 $\pm 4 \pm 4$	25k	<sup>10</sup> BIRD	89 LASS	—	$11 K^- p \rightarrow \bar{K}^0 \pi^- p$
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## NEUTRAL ONLY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$109 \pm 5$	OUR AVERAGE	Error includes scale factor of 1.9. See the ideogram below.		
116.5 $\pm 3.6 \pm 1.7$		<sup>11</sup> ASTON	88 LASS	$11 K^- p \rightarrow K^- \pi^+ n$
129 $\pm 15 \pm 15$		<sup>11</sup> ASTON	87 LASS	$11 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
131 $\pm 24 \pm 20$		<sup>11</sup> ASTON	84B LASS	$11 K^- p \rightarrow \bar{K}^0 2\pi n$
143 $\pm 34$		<sup>11</sup> BAUBILLIER	82B HBC	$8.25 K^- p \rightarrow N K_S^0 \pi \pi$
98 $\pm 8$		<sup>11</sup> ASTON	81C LASS	$11 K^- p \rightarrow K^- \pi^+ n$
140 $\pm 30$		<sup>11</sup> ETKIN	80 SPEC	$6 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
98 $\pm 5$		<sup>11</sup> ESTABROOKS	78 ASPK	$13 K^\pm p \rightarrow p K \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
113.7 $\pm 9.2$	1786 $\pm 127$	<sup>12</sup> AUBERT	07AK BABR	$10.6 e^+ e^- \rightarrow K^{*0} K^\pm \pi^\mp \gamma$
125 $\pm 29$	300	<sup>8</sup> HENDRICK	76 DBC	$8.25 K^+ N \rightarrow K^+ \pi N$
116 $\pm 18$	800	MCCUBBIN	75 HBC	$3.6 K^- p \rightarrow K^- \pi^+ n$
61 $\pm 14$		<sup>13</sup> LINGLIN	73 HBC	$2-13 K^+ p \rightarrow K^+ \pi^- X$
116.6 $\pm 10.3$ —15.5	1800	AGUILAR-...	71B HBC	$3.9, 4.6 K^- p$
144 $\pm 24.0$	600	<sup>8</sup> CORDS	71 DBC	$9 K^+ n \rightarrow K^+ \pi^- p$
101 $\pm 10$	2200	DAVIS	69 HBC	$12 K^+ p \rightarrow K^+ \pi^- \pi^+ p$

WEIGHTED AVERAGE  
109 $\pm$ 5 (Error scaled by 1.9)



$K_2^*(1430)^0$  width (MeV)

<sup>8</sup> Errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>9</sup> Number of events in peak re-evaluated by us.

<sup>10</sup> From a partial wave amplitude analysis.

<sup>11</sup> From phase shift or partial-wave analysis.

<sup>12</sup> Systematic errors not estimated.

<sup>13</sup> From pole extrapolation, using world  $K^+ p$  data summary tape.

 $K_2^*(1430)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $K\pi$	(49.9 $\pm$ 1.2) %	
$\Gamma_2$ $K^*(892)\pi$	(24.7 $\pm$ 1.5) %	
$\Gamma_3$ $K^*(892)\pi\pi$	(13.4 $\pm$ 2.2) %	
$\Gamma_4$ $K\rho$	( 8.7 $\pm$ 0.8) %	S=1.2

# Meson Particle Listings

## $K_2^*(1430)$

$\Gamma_5$	$K\omega$	$(2.9 \pm 0.8) \%$	
$\Gamma_6$	$K^+\gamma$	$(2.4 \pm 0.5) \times 10^{-3}$	$S=1.1$
$\Gamma_7$	$K\eta$	$(1.5^{+3.4}_{-1.0}) \times 10^{-3}$	$S=1.3$
$\Gamma_8$	$K\omega\pi$	$< 7.2 \times 10^{-4}$	$CL=95\%$
$\Gamma_9$	$K^0\gamma$	$< 9 \times 10^{-4}$	$CL=90\%$

### CONSTRAINED FIT INFORMATION

An overall fit to the total width, a partial width, and 10 branching ratios uses 31 measurements and one constraint to determine 8 parameters. The overall fit has a  $\chi^2 = 20.2$  for 24 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−9						
$x_3$	−40	−73					
$x_4$	−8	36	−52				
$x_5$	−11	−3	−26	−7			
$x_6$	−1	−1	−1	−1	0		
$x_7$	−4	−7	−5	−5	−2	0	
$\Gamma$	0	0	0	0	0	−13	0
	$x_1$	$x_2$	$x_3$	$x_4$	$x_5$	$x_6$	$x_7$

Mode	Rate (MeV)	Scale factor
$\Gamma_1$ $K\pi$	$49.1 \pm 1.8$	
$\Gamma_2$ $K^*(892)\pi$	$24.3 \pm 1.6$	
$\Gamma_3$ $K^*(892)\pi\pi$	$13.2 \pm 2.2$	
$\Gamma_4$ $K\rho$	$8.5 \pm 0.8$	1.2
$\Gamma_5$ $K\omega$	$2.9 \pm 0.8$	
$\Gamma_6$ $K^+\gamma$	$0.24 \pm 0.05$	1.1
$\Gamma_7$ $K\eta$	$0.15^{+0.33}_{-0.10}$	1.3

### $K_2^*(1430)$ PARTIAL WIDTHS

$\Gamma(K^+\gamma)$							$\Gamma_6$
VALUE (keV)		DOCUMENT ID	TECN	CHG	COMMENT		
<b>241 ± 50 OUR FIT</b>		Error includes scale factor of 1.1.					
<b>240 ± 45</b>		CIHANGIR	82	SPEC	+	200 $K^+Z \rightarrow ZK^+\pi^0$ , $ZK_S^0\pi^+$	

$\Gamma(K^0\gamma)$							$\Gamma_9$
VALUE (keV)	CL%	DOCUMENT ID	TECN	CHG	COMMENT		
< 5.4	90	ALAVI-HARATI02b	KTEV		$K + A \rightarrow K^* + A$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
< 84	90	CARLSMITH	87	SPEC	0	60–200 $K_L^0 A \rightarrow K_S^0 \pi^0 A$	

### $K_2^*(1430)$ BRANCHING RATIOS

$\Gamma(K\pi) / \Gamma_{\text{total}}$							$\Gamma_1 / \Gamma$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.499 ± 0.012 OUR FIT</b>							
<b>0.488 ± 0.014 OUR AVERAGE</b>							
$0.485 \pm 0.006 \pm 0.020$	14	ASTON	88	LASS	0	11 $K^-p \rightarrow K^- \pi^+ n$	
$0.49 \pm 0.02$	14	ESTABROOKS	78	ASPK	±	13 $K^\pm p \rightarrow p K \pi$	

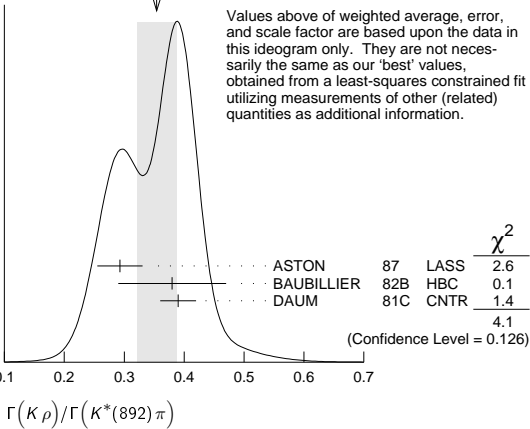
$\Gamma(K^*(892)\pi) / \Gamma(K\pi)$							$\Gamma_2 / \Gamma_1$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.496 ± 0.034 OUR FIT</b>							
<b>0.47 ± 0.04 OUR AVERAGE</b>							
$0.44 \pm 0.09$		ASTON	84b	LASS	0	11 $K^-p \rightarrow \bar{K}^0 2\pi n$	
$0.62 \pm 0.19$		LAUSCHER	75	HBC	0	10, 16 $K^-p \rightarrow K^- \pi^+ n$	
$0.54 \pm 0.16$		DEHM	74	DBC	0	4.6 $K^+ N$	
$0.47 \pm 0.08$		AGUILAR-...	71b	HBC		3.9, 4.6 $K^- p$	
$0.47 \pm 0.10$		BASSANO	67	HBC	−0	4.6, 5.0 $K^- p$	
$0.45 \pm 0.13$		BADIER	65c	HBC	−	3 $K^- p$	

$\Gamma(K\omega) / \Gamma(K\pi)$							$\Gamma_5 / \Gamma_1$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.059 ± 0.017 OUR FIT</b>							
<b>0.070 ± 0.035 OUR AVERAGE</b>							
$0.05 \pm 0.04$		AGUILAR-...	71b	HBC		3.9, 4.6 $K^- p$	
$0.13 \pm 0.07$		BASSOMPIE-...	69	HBC	0	5 $K^+ p$	

$\Gamma(K\rho) / \Gamma(K\pi)$							$\Gamma_4 / \Gamma_1$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.174 ± 0.017 OUR FIT</b>		Error includes scale factor of 1.2.					
<b>0.150 ± 0.029 OUR AVERAGE</b>							
$0.18 \pm 0.05$		ASTON	84b	LASS	0	11 $K^-p \rightarrow \bar{K}^0 2\pi n$	
$0.02^{+0.10}_{-0.02}$		DEHM	74	DBC	0	4.6 $K^+ N$	
$0.16 \pm 0.05$		AGUILAR-...	71b	HBC		3.9, 4.6 $K^- p$	
$0.14 \pm 0.10$		BASSANO	67	HBC	−0	4.6, 5.0 $K^- p$	
$0.14 \pm 0.07$		BADIER	65c	HBC	−	3 $K^- p$	

$\Gamma(K\rho) / \Gamma(K^*(892)\pi)$							$\Gamma_4 / \Gamma_2$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.350 ± 0.031 OUR FIT</b>		Error includes scale factor of 1.4.					
<b>0.354 ± 0.033 OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.					
$0.293 \pm 0.032 \pm 0.020$		ASTON	87	LASS	0	11 $K^-p \rightarrow \bar{K}^0 \pi^+ \pi^- n$	
$0.38 \pm 0.09$		BAUBILLIER	82b	HBC	0	8.25 $K^-p \rightarrow NK_S^0 \pi \pi$	
$0.39 \pm 0.03$		DAUM	81c	CNTR		63 $K^-p \rightarrow K^- 2\pi p$	

WEIGHTED AVERAGE  
0.354 ± 0.033 (Error scaled by 1.4)



$\Gamma(K\omega) / \Gamma(K^*(892)\pi)$							$\Gamma_5 / \Gamma_2$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.118 ± 0.034 OUR FIT</b>							
<b>0.10 ± 0.04</b>		FIELD	67	HBC	−	3.8 $K^- p$	

$\Gamma(K\eta) / \Gamma(K^*(892)\pi)$							$\Gamma_7 / \Gamma_2$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.006 ± 0.014 OUR FIT</b>		Error includes scale factor of 1.2.					
<b>0.07 ± 0.04</b>		FIELD	67	HBC	−	3.8 $K^- p$	

$\Gamma(K\eta) / \Gamma(K\pi)$							$\Gamma_7 / \Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.0030 ± 0.0070 OUR FIT</b>		Error includes scale factor of 1.3.					
<b>0 ± 0.0056</b>		15 ASTON	88b	LASS	−	11 $K^-p \rightarrow K^- \eta p$	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
< 0.04	95	AGUILAR-...	71b	HBC		3.9, 4.6 $K^- p$	
< 0.065		16 BASSOMPIE-...	69	HBC		5.0 $K^+ p$	
< 0.02		BISHOP	69	HBC		3.5 $K^+ p$	

$\Gamma(K^*(892)\pi\pi) / \Gamma_{\text{total}}$							$\Gamma_3 / \Gamma$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.134 ± 0.022 OUR FIT</b>							
<b>0.12 ± 0.04</b>		17 GOLDBERG	76	HBC	−	3 $K^-p \rightarrow p \bar{K}^0 \pi \pi$	

$\Gamma(K^*(892)\pi\pi) / \Gamma(K\pi)$							$\Gamma_3 / \Gamma_1$
VALUE		DOCUMENT ID	TECN	CHG	COMMENT		
<b>0.27 ± 0.05 OUR FIT</b>							
<b>0.21 ± 0.08</b>		16, 17 JONGEJANS	78	HBC	−	4 $K^-p \rightarrow p \bar{K}^0 \pi \pi$	

$\Gamma(K\omega\pi) / \Gamma_{\text{total}}$							$\Gamma_8 / \Gamma$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>&lt; 0.72</b>	95	0	JONGEJANS	78	HBC	4 $K^-p \rightarrow p \bar{K}^0 4\pi$	

14 From phase shift analysis.  
15 ASTON 88b quote < 0.0092 at CL=95%. We convert this to a central value and 1 sigma error in order to be able to use it in our constrained fit.  
16 Restated by us.  
17 Assuming  $\pi\pi$  system has isospin 1, which is supported by the data.

See key on page 601

Meson Particle Listings

$K_2^*(1430)$ ,  $K(1460)$ ,  $K_2(1580)$ ,  $K(1630)$

$K_2^*(1430)$  REFERENCES

AUBERT	07AK	PR D76 012008	B. Aubert <i>et al.</i>	(BABAR Collab.)
ALAVI-HARATI	02B	PRL 89 072001	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
BIRD	89	SLAC-332	P.F. Bird	(SLAC)
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	88B	PL B201 169	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	87	NP B292 693	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
CARLSMITH	87	PR D36 3502	D. Carlsmith <i>et al.</i>	(EFI, SACL)
ASTON	84B	NP B247 261	D. Aston <i>et al.</i>	(SLAC, CARL, OTTA)
BAUBILLIER	84B	ZPHY C26 37	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
BAUBILLIER	82B	NP B202 21	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
CHANGIR	82	PL 117B 123	S. Chhangir <i>et al.</i>	(FNAL, MINN, ROCH)
CLELAND	82	NP B208 189	W.E. Cleland <i>et al.</i>	(DURH, GEVA, LAUS+)
ASTON	81C	PL 106B 235	D. Aston <i>et al.</i>	(SLAC, CARL, OTTA) JP
DAUM	81C	NP B187 1	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
TOAFF	81	PR D23 1500	S. Toaff <i>et al.</i>	(ANL, KANS)
ETKIN	80	PR D22 42	A. Etkin <i>et al.</i>	(BNL, CUNY) JP
ESTABROOKS	78	NP B133 490	P.G. Estabrooks <i>et al.</i>	(MCGI, CARL, DURH+)
Also		PR D17 658	P.G. Estabrooks <i>et al.</i>	(MCGI, CARL, DURH+)
JONGEJANS	78	NP B139 383	B. Jongejans <i>et al.</i>	(ZEEM, CERN, NIJ+)
MARTIN	78	NP B134 392	A.D. Martin <i>et al.</i>	(DURH, GEVA)
BOWLER	77	NP B126 31	M.G. Bowler <i>et al.</i>	(OXF)
GOLDBERG	76	LNC 17 253	J. Goldberg	(HAIF)
HENDRICK	76	NP B112 189	K. Hendrickx <i>et al.</i>	(MONS, SACL, PARIS+)
LAUSCHER	75	NP B86 189	P. Lauscher <i>et al.</i>	(ABCLV Collab.) JP
MCCUBBIN	75	NP B86 13	N.A. McCubbin, L. Lyons	(OXF)
DEHM	74	NP B75 47	G. Dehm <i>et al.</i>	(MPIM, BRUX, MONS, CERN)
LINGLIN	73	NP B55 408	D. Linglin	(CERN)
AGUILAR...	71B	PR D4 2583	M. Aguilar-Benitez, R.L. Eisner, J.B. Kinson	(BNL)
BARNHAM	71C	NP B28 171	K.W.J. Barnham <i>et al.</i>	(BIRM, GLAS)
CORDS	71	PR D4 1974	D. Cords <i>et al.</i>	(PURD, UCD, IUPU)
BASSOMPIERRE...	69	NP B13 189	G. Bassompierre <i>et al.</i>	(CERN, BRUX) JP
BISHOP	69	NP B9 403	J.M. Bishop <i>et al.</i>	(WISC)
CRENNELL	69D	PRL 22 487	D.J. Crennell <i>et al.</i>	(BNL)
DAVIS	69	PRL 23 1071	P.J. Davis <i>et al.</i>	(LRL)
LIND	69	NP B14 1	V.G. Lind <i>et al.</i>	(LRL) JP
SCHWEING...	68	PR 166 1317	F. Schweingruber <i>et al.</i>	(ANL, NWES)
Also			F.L. Schweingruber	(NWES, NWES)
BASSANO	67	PRL 19 968	D. Bassano <i>et al.</i>	(BNL, SYRA)
FIELD	67	PL 24B 638	J.H. Field <i>et al.</i>	(UCSD)
BADIER	65C	PL 19 612	J. Badier <i>et al.</i>	(EPOL, SACL, AMST)

$K(1460)$

$I(J^P) = \frac{1}{2}(0^-)$

OMITTED FROM SUMMARY TABLE  
Observed in  $K\pi\pi$  partial-wave analysis.

$K(1460)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
...	...	...	...	...
~ 1460	DAUM	81c	CNTR	— 63 $K^-p \rightarrow K^-2\pi p$
~ 1400	<sup>1</sup> BRANDENB...	76B	ASPK	± 13 $K^\pm p \rightarrow K^\pm 2\pi p$
<sup>1</sup> Coupled mainly to $K\bar{f}_0(1370)$ . Decay into $K^*(892)\pi$ seen.				

$K(1460)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
...	...	...	...	...
~ 260	DAUM	81c	CNTR	— 63 $K^-p \rightarrow K^-2\pi p$
~ 250	<sup>2</sup> BRANDENB...	76B	ASPK	± 13 $K^\pm p \rightarrow K^\pm 2\pi p$
<sup>2</sup> Coupled mainly to $K\bar{f}_0(1370)$ . Decay into $K^*(892)\pi$ seen.				

$K(1460)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K^*(892)\pi$	seen
$\Gamma_2$ $K\rho$	seen
$\Gamma_3$ $K_0^*(1430)\pi$	seen

$K(1460)$  PARTIAL WIDTHS

$\Gamma(K^*(892)\pi)$				$\Gamma_1$
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
...	...	...	...	...
~ 109	DAUM	81c	CNTR	63 $K^-p \rightarrow K^-2\pi p$
$\Gamma(K\rho)$				$\Gamma_2$
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
...	...	...	...	...
~ 34	DAUM	81c	CNTR	63 $K^-p \rightarrow K^-2\pi p$
$\Gamma(K_0^*(1430)\pi)$				$\Gamma_3$
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
...	...	...	...	...
~ 117	DAUM	81c	CNTR	63 $K^-p \rightarrow K^-2\pi p$

$K(1460)$  REFERENCES

DAUM	81C	NP B187 1	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
BRANDENB...	76B	PRL 36 1239	G.W. Brandenburg <i>et al.</i>	(SLAC) JP

$K_2(1580)$

$I(J^P) = \frac{1}{2}(2^-)$

OMITTED FROM SUMMARY TABLE  
Seen in partial-wave analysis of the  $K^-\pi^+\pi^-$  system. Needs confirmation.

$K_2(1580)$  MASS

VALUE (MeV)	DOCUMENT ID	CHG	COMMENT
...	...	...	...
~ 1580	OTTER	79	— 10,14,16 $K^-p$

$K_2(1580)$  WIDTH

VALUE (MeV)	DOCUMENT ID	CHG	COMMENT
...	...	...	...
~ 110	OTTER	79	— 10,14,16 $K^-p$

$K_2(1580)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K^*(892)\pi$	seen
$\Gamma_2$ $K_2^*(1430)\pi$	possibly seen

$K_2(1580)$  BRANCHING RATIOS

$\Gamma(K^*(892)\pi)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
seen	OTTER	79	HBC	— 10,14,16 $K^-p$	
...	...	...	...	...	...
possibly seen	GULER	11	BELL	$B^+ \rightarrow J/\psi K^+\pi^+\pi^-$	

$\Gamma(K_2^*(1430)\pi)/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
possibly seen	OTTER	79	HBC	— 10,14,16 $K^-p$	

$K_2(1580)$  REFERENCES

GULER	11	PR D83 032005	H. Guler <i>et al.</i>	(BELLE Collab.)
OTTER	79	NP B147 1	G. Otter <i>et al.</i>	(AACH3, BERL, CERN, LOIC+)

$K(1630)$

$I(J^P) = \frac{1}{2}(?^?)$

OMITTED FROM SUMMARY TABLE  
Seen as a narrow peak, compatible with the experimental resolution, in the invariant mass of the  $K_S^0\pi^+\pi^-$  system produced in  $\pi^-p$  interactions at high momentum transfers.

$K(1630)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$1629 \pm 7$	~ 75	KARNAUKHOV98	BC	$16.0 \pi^-p \rightarrow (K_S^0\pi^+\pi^-) X^+\pi^-X^0$

$K(1630)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$16^{+19}_{-16}$	~ 75	<sup>1</sup> KARNAUKHOV98	BC	$16.0 \pi^-p \rightarrow (K_S^0\pi^+\pi^-) X^+\pi^-X^0$

<sup>1</sup> Compatible with an experimental resolution of  $14 \pm 1$  MeV.

$K(1630)$  DECAY MODES

Mode
$\Gamma_1$ $K_S^0\pi^+\pi^-$

$K(1630)$  REFERENCES

KARNAUKHOV	98	PAN 61 203	V.M. Karnaukhov, C. Coca, V.I. Moroz
Translated from YAF 61 252.			

Meson Particle Listings

$K_1(1650)$ ,  $K^*(1680)$ ,  $K_2(1770)$

$K_1(1650)$

$I(J^P) = \frac{1}{2}(1^+)$

OMITTED FROM SUMMARY TABLE

This entry contains various peaks in strange meson systems ( $K^+\phi$ ,  $K\pi\pi$ ) reported in partial-wave analysis in the 1600–1900 mass region.

$K_1(1650)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>1650±50</b>	FRAME	86	OMEG +	13 $K^+p \rightarrow \phi K^+p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 1840	ARMSTRONG	83	OMEG –	18.5 $K^-p \rightarrow 3Kp$
~ 1800	DAUM	81c	CNTR –	63 $K^-p \rightarrow K^-2\pi p$

$K_1(1650)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>150±50</b>	FRAME	86	OMEG +	13 $K^+p \rightarrow \phi K^+p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
~ 250	DAUM	81c	CNTR –	63 $K^-p \rightarrow K^-2\pi p$

$K_1(1650)$  DECAY MODES

Mode	
$\Gamma_1$	$K\pi\pi$
$\Gamma_2$	$K\phi$

$K_1(1650)$  REFERENCES

FRAME	86	NP B276 667	D. Frame <i>et al.</i>	(GLAS)
ARMSTRONG	83	NP B221 1	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
DAUM	81c	NP B187 1	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)

$K^*(1680)$

$I(J^P) = \frac{1}{2}(1^-)$

$K^*(1680)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>1717±27 OUR AVERAGE</b>	Error includes scale factor of 1.4.			
1677±10±32	ASTON	88	LASS 0	11 $K^-p \rightarrow K^-\pi^+n$
1735±10±20	ASTON	87	LASS 0	11 $K^-p \rightarrow \overline{K}^0\pi^+\pi^-n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1678±64	BIRD	89	LASS –	11 $K^-p \rightarrow \overline{K}^0\pi^-\pi^-p$
1800±70	ETKIN	80	MPS 0	6 $K^-p \rightarrow \overline{K}^0\pi^+\pi^-n$
~ 1650	ESTABROOKS	78	ASPK 0	13 $K^\pm p \rightarrow K^\pm\pi^\pm n$

$K^*(1680)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>322±110 OUR AVERAGE</b>	Error includes scale factor of 4.2.			
205± 16±34	ASTON	88	LASS 0	11 $K^-p \rightarrow K^-\pi^+n$
423± 18±30	ASTON	87	LASS 0	11 $K^-p \rightarrow \overline{K}^0\pi^+\pi^-n$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
454±270	BIRD	89	LASS –	11 $K^-p \rightarrow \overline{K}^0\pi^-\pi^-p$
170± 30	ETKIN	80	MPS 0	6 $K^-p \rightarrow \overline{K}^0\pi^+\pi^-n$
250 to 300	ESTABROOKS	78	ASPK 0	13 $K^\pm p \rightarrow K^\pm\pi^\pm n$

$K^*(1680)$  DECAY MODES

Mode		Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$K\pi$	(38.7±2.5) %
$\Gamma_2$	$K\rho$	(31.4 <sup>+5.0</sup> <sub>–2.1</sub> ) %
$\Gamma_3$	$K^*(892)\pi$	(29.9 <sup>+2.2</sup> <sub>–5.0</sub> ) %

CONSTRAINED FIT INFORMATION

An overall fit to 4 branching ratios uses 4 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 2.9$  for 2 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle\delta x_i\delta x_j\rangle/(\delta x_i\delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	–36	
$x_3$	–39	–72
	$x_1$	$x_2$

$K^*(1680)$  BRANCHING RATIOS

$\Gamma(K\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_1/\Gamma$
<b>0.387±0.026 OUR FIT</b>					
<b>0.388±0.014±0.022</b>	ASTON	88	LASS 0	11 $K^-p \rightarrow K^-\pi^+n$	

$\Gamma(K\rho)/\Gamma(K^*(892)\pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_1/\Gamma_3$
<b>1.30<sup>+0.23</sup><sub>–0.14</sub> OUR FIT</b>					
<b>2.8 ±1.1</b>	ASTON	84	LASS 0	11 $K^-p \rightarrow \overline{K}^02\pi n$	

$\Gamma(K\rho)/\Gamma(K\pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_2/\Gamma_1$
<b>0.81<sup>+0.14</sup><sub>–0.09</sub> OUR FIT</b>					
<b>1.2 ±0.4</b>	ASTON	84	LASS 0	11 $K^-p \rightarrow \overline{K}^02\pi n$	

$\Gamma(K\rho)/\Gamma(K^*(892)\pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_2/\Gamma_3$
<b>1.05<sup>+0.27</sup><sub>–0.11</sub> OUR FIT</b>					
<b>0.97±0.09±0.30 0.10</b>	ASTON	87	LASS 0	11 $K^-p \rightarrow \overline{K}^0\pi^+\pi^-n$	

$K^*(1680)$  REFERENCES

BIRD	89	SLAC-332	P.F. Bird	(SLAC)
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	87	NP B292 693	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	84	PL 149B 258	D. Aston <i>et al.</i>	(SLAC, CARL, OTTA)JP
ETKIN	80	PR D22 42	A. Etkin <i>et al.</i>	(BNL, CUNY)JP
ESTABROOKS	78	NP B133 490	P.G. Estabrooks <i>et al.</i>	(MCGI, CARL, DURH+)JP

$K_2(1770)$

$I(J^P) = \frac{1}{2}(2^-)$

See our mini-review in the 2004 edition of this *Review*, PDG 04.

$K_2(1770)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1773± 8</b>	<sup>1</sup>	ASTON	93	LASS	11 $K^-p \rightarrow K^-\omega p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1743±15		TIKHOMIROV	03	SPEC	40.0 $\pi^-\overline{C} \rightarrow K_S^0 K_S^0 K_L^0 X$
1810±20		FRAME	86	OMEG +	13 $K^+p \rightarrow \phi K^+p$
~ 1730		ARMSTRONG	83	OMEG –	18.5 $K^-p \rightarrow 3Kp$
~ 1780		<sup>2</sup> DAUM	81c	CNTR –	63 $K^-p \rightarrow K^-2\pi p$
1710±15	60	CHUNG	74	HBC –	7.3 $K^-p \rightarrow K^-\omega p$
1767± 6		BLIEDEN	72	MMS –	11–16 $K^-p$
1730±20	306	<sup>3</sup> FIRESTONE	72b	DBC +	12 $K^+d$
1765±40		<sup>4</sup> COLLEY	71	HBC +	10 $K^+p \rightarrow K2\pi N$
1740		DENEGRI	71	DBC –	12.6 $K^-d \rightarrow \overline{K}2\pi d$
1745±20		AGUILAR...	70c	HBC –	4.6 $K^-p$
1780±15		BARTSCH	70c	HBC –	10.1 $K^-p$
1760±15		LUDLAM	70	HBC –	12.6 $K^-p$

<sup>1</sup> From a partial wave analysis of the  $K^-\omega$  system.  
<sup>2</sup> From a partial wave analysis of the  $K^-2\pi$  system.  
<sup>3</sup> Produced in conjunction with excited deuteron.  
<sup>4</sup> Systematic errors added correspond to spread of different fits.

$K_2(1770)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>186±14</b>	<sup>5</sup>	ASTON	93	LASS	11 $K^-p \rightarrow K^-\omega p$

See key on page 601

Meson Particle Listings

$K_2(1770)$ ,  $K_3^*(1780)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$147 \pm 70$		TIKHOMIROV	03	SPEC	$40.0 \frac{\pi^- C}{K_S^0 K_S^0 K_L^0 X}$
$140 \pm 40$		FRAME	86	OMEG +	$13 K^+ p \rightarrow \phi K^+ p$
$\sim 220$		ARMSTRONG	83	OMEG -	$18.5 K^- p \rightarrow 3 K p$
$\sim 210$		6 DAUM	81c	CNTR -	$63 K^- p \rightarrow K^- 2\pi p$
$110 \pm 50$	60	CHUNG	74	HBC -	$7.3 K^- p \rightarrow K^- \omega p$
$100 \pm 26$		BLIEDEN	72	MMS -	$11-16 K^- p$
$210 \pm 30$	306	7 FIRESTONE	72B	DBC +	$12 K^+ d$
$90 \pm 70$		8 COLLEY	71	HBC +	$10 K^+ p \rightarrow K 2\pi N$
130		DENEGRI	71	DBC -	$12.6 K^- d \rightarrow \bar{K} 2\pi d$
$100 \pm 50$		AGUILAR-...	70c	HBC -	$4.6 K^- p$
$138 \pm 40$		BARTSCH	70c	HBC -	$10.1 K^- p$
$50 \pm 40$ $-20$		LUDLAM	70	HBC -	$12.6 K^- p$

<sup>5</sup> From a partial wave analysis of the  $K^- \omega$  system.

<sup>6</sup> From a partial wave analysis of the  $K^- 2\pi$  system.

<sup>7</sup> Produced in conjunction with excited deuteron.

<sup>8</sup> Systematic errors added correspond to spread of different fits.

$K_2(1770)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad K \pi \pi$	
$\Gamma_2 \quad K_2^*(1430) \pi$	dominant
$\Gamma_3 \quad K^*(892) \pi$	seen
$\Gamma_4 \quad K f_2(1270)$	seen
$\Gamma_5 \quad K f_0(980)$	
$\Gamma_6 \quad K \phi$	seen
$\Gamma_7 \quad K \omega$	seen

$K_2(1770)$  BRANCHING RATIOS

$\Gamma(K_2^*(1430)\pi)/\Gamma(K\pi\pi)$   $\Gamma_2/\Gamma_1$   
( $K_2^*(1430) \rightarrow K\pi$ )

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 0.03$	DAUM	81c	CNTR	$63 K^- p \rightarrow K^- 2\pi p$
$\sim 1.0$	9 FIRESTONE	72B	DBC +	$12 K^+ d$
$<1.0$	COLLEY	71	HBC	$10 K^+ p$
$0.2 \pm 0.2$	AGUILAR-...	70c	HBC -	$4.6 K^- p$
$<1.0$	BARTSCH	70c	HBC -	$10.1 K^- p$
1.0	BARBARO-...	69	HBC +	$12.0 K^+ p$

<sup>9</sup> Produced in conjunction with excited deuteron.

$\Gamma(K^*(892)\pi)/\Gamma(K\pi\pi)$   $\Gamma_3/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$\sim 0.23$	DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$

$\Gamma(K f_2(1270))/\Gamma(K\pi\pi)$   $\Gamma_4/\Gamma_1$   
( $f_2(1270) \rightarrow \pi\pi$ )

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$\sim 0.74$	DAUM	81c	CNTR $63 K^- p \rightarrow K^- 2\pi p$

$\Gamma(K f_0(980))/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
possibly seen	TIKHOMIROV	03	SPEC $40.0 \frac{\pi^- C}{K_S^0 K_S^0 K_L^0 X}$

$\Gamma(K\phi)/\Gamma_{\text{total}}$   $\Gamma_6/\Gamma$

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	ARMSTRONG	83	OMEG -	$18.5 K^- p \rightarrow K^- \phi N$

$\Gamma(K\omega)/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
seen	OTTER	81	HBC $\pm$	$8.25, 10, 16 K^\pm p$
seen	CHUNG	74	HBC -	$7.3 K^- p \rightarrow K^- \omega p$

$K_2(1770)$  REFERENCES

PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
TIKHOMIROV	03	PAN 66 828	G.D. Tikhomirov <i>et al.</i>	
ASTON	93	PL B308 186	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
FRAME	86	NP B276 667	D. Frame <i>et al.</i>	(GLAS)
ARMSTRONG	83	NP B221 1	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
DAUM	81c	NP B187 1	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIN+)
OTTER	81	NP B181 1	G. Otter	(AACH3, BERL, LOIC, VIEN, BIRM+)
CHUNG	74	PL 518 413	S.U. Chung <i>et al.</i>	(BNL)
BLIEDEN	72	PL 398 668	H.R. Blieden <i>et al.</i>	(STON, NEAS)
FIRESTONE	72B	PR D5 505	A. Firestone <i>et al.</i>	(LBL)
COLLEY	71	NP B26 71	D.C. Colley <i>et al.</i>	(BIRM, GLAS)
DENEGRI	71	NP B28 13	D. Denegri <i>et al.</i>	(JHU)JP
AGUILAR-...	70c	PRL 25 54	M. Aguilar-Benítez <i>et al.</i>	(BNL)
BARTSCH	70c	PL 33B 186	J. Bartsch <i>et al.</i>	(AACH, BERL, CERN+)
LUDLAM	70	PR D2 1234	T. Ludlam, J. Sandweiss, A.J. Slaughter	(YALE)
BARBARO-...	69	PRL 22 1207	A. Barbaro-Gallieri <i>et al.</i>	(LRL)

$K_3^*(1780)$

$I(J^P) = \frac{1}{2}(3^-)$

$K_3^*(1780)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1776 ± 7 OUR AVERAGE</b>		Error includes scale factor of 1.1.			
$1781 \pm 8 \pm 4$		1 ASTON	88	LASS	0 $11 K^- p \rightarrow K^- \pi^+ n$
$1740 \pm 14 \pm 15$		1 ASTON	87	LASS	0 $11 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
$1779 \pm 11$		2 BALDI	76	SPEC +	10 $K^+ p \rightarrow K^0 \pi^+ p$
$1776 \pm 26$		3 BRANDENB...	76D	ASPK	0 $13 K^\pm p \rightarrow K^\pm \pi^\mp N$
$1720 \pm 10 \pm 15$	6111	4 BIRD	89	LASS -	11 $K^- p \rightarrow \bar{K}^0 \pi^- p$
$1749 \pm 10$		ASTON	88B	LASS -	11 $K^- p \rightarrow K^- \eta p$
$1780 \pm 9$	300	BAUBILLIER	84B	HBC -	8.25 $K^- p \rightarrow \bar{K}^0 \pi^- p$
$1790 \pm 15$		BAUBILLIER	82B	HBC	0 $8.25 K^- p \rightarrow K_S^0 2\pi N$
$1784 \pm 9$	2060	CLELAND	82	SPEC $\pm$	50 $K^+ p \rightarrow K_S^0 \pi^\pm p$
$1786 \pm 15$		5 ASTON	81D	LASS	0 $11 K^- p \rightarrow K^- \pi^+ n$
$1762 \pm 9$	190	TOAFF	81	HBC -	6.5 $K^- p \rightarrow \bar{K}^0 \pi^- p$
$1850 \pm 50$		ETKIN	80	MPS	0 $6 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^-$
$1812 \pm 28$		BEUSCH	78	OMEG	10 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
$1786 \pm 8$		CHUNG	78	MPS	0 $6 K^- p \rightarrow K^- \pi^+ n$

<sup>1</sup> From energy-independent partial-wave analysis.

<sup>2</sup> From a fit to  $Y_6^2$  moment.  $J^P = 3^-$  found.

<sup>3</sup> Confirmed by phase shift analysis of ESTABROOKS 78, yields  $J^P = 3^-$ .

<sup>4</sup> From a partial wave amplitude analysis.

<sup>5</sup> From a fit to the  $Y_6^0$  moment.

$K_3^*(1780)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>159 ± 21 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.			
$203 \pm 30 \pm 8$		6 ASTON	88	LASS	0 $11 K^- p \rightarrow K^- \pi^+ n$
$171 \pm 42 \pm 20$		6 ASTON	87	LASS	0 $11 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
$135 \pm 22$		7 BALDI	76	SPEC +	10 $K^+ p \rightarrow K^0 \pi^+ p$
$187 \pm 31 \pm 20$	6111	8 BIRD	89	LASS -	11 $K^- p \rightarrow \bar{K}^0 \pi^- p$
$193 \pm 51$ $-37$		ASTON	88B	LASS -	11 $K^- p \rightarrow K^- \eta p$
$99 \pm 30$	300	BAUBILLIER	84B	HBC -	8.25 $K^- p \rightarrow \bar{K}^0 \pi^- p$
$\sim 130$		BAUBILLIER	82B	HBC	0 $8.25 K^- p \rightarrow K_S^0 2\pi N$
$191 \pm 24$	2060	CLELAND	82	SPEC $\pm$	50 $K^+ p \rightarrow K_S^0 \pi^\pm p$
$225 \pm 60$		9 ASTON	81D	LASS	0 $11 K^- p \rightarrow K^- \pi^+ n$
$\sim 80$	190	TOAFF	81	HBC -	6.5 $K^- p \rightarrow \bar{K}^0 \pi^- p$
$240 \pm 50$		ETKIN	80	MPS	0 $6 K^- p \rightarrow \bar{K}^0 \pi^+ \pi^-$
$181 \pm 44$		10 BEUSCH	78	OMEG	10 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
$96 \pm 31$		CHUNG	78	MPS	0 $6 K^- p \rightarrow K^- \pi^+ n$
$270 \pm 70$		11 BRANDENB...	76D	ASPK	0 $13 K^\pm p \rightarrow K^\pm \pi^\mp N$

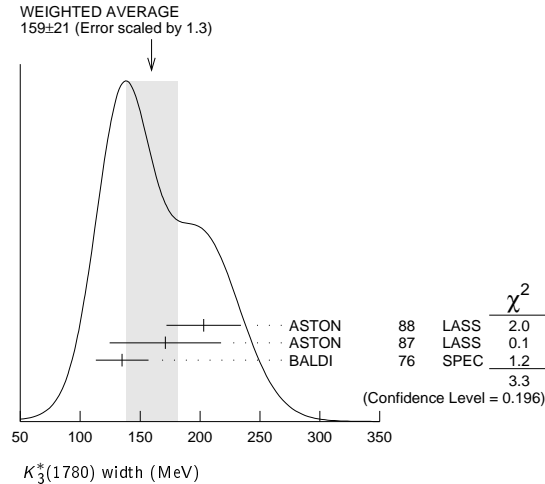
<sup>6</sup> From energy-independent partial-wave analysis.

<sup>7</sup> From a fit to  $Y_6^2$  moment.  $J^P = 3^-$  found.

Meson Particle Listings

$K_3^*(1780)$ ,  $K_2(1820)$

<sup>8</sup> From a partial wave amplitude analysis.  
<sup>9</sup> From a fit to  $Y_6^0$  moment.  
<sup>10</sup> Errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.  
<sup>11</sup> ESTABROOKS 78 find that BRANDENBURG 76D data are consistent with 175 MeV width. Not averaged.



$K_3^*(1780)$ DECAY MODES				
Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level		
$\Gamma_1$ $K\rho$	(31 $\pm$ 9 ) %			
$\Gamma_2$ $K^*(892)\pi$	(20 $\pm$ 5 ) %			
$\Gamma_3$ $K\pi$	(18.8 $\pm$ 1.0) %			
$\Gamma_4$ $K\eta$	(30 $\pm$ 13 ) %			
$\Gamma_5$ $K_2^*(1430)\pi$	< 16 %	95%		

**CONSTRAINED FIT INFORMATION**

An overall fit to 3 branching ratios uses 4 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 0.0$  for 1 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle\delta x_i\delta x_j\rangle/(\delta x_i\delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	85		
$x_3$	18	21	
$x_4$	−98	−94	−27
	$x_1$	$x_2$	$x_3$

$K_3^*(1780)$ BRANCHING RATIOS					
$\Gamma(K\rho)/\Gamma(K^*(892)\pi)$					$\Gamma_1/\Gamma_2$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>1.52<math>\pm</math>0.23 OUR FIT</b>					
<b>1.52<math>\pm</math>0.21<math>\pm</math>0.10</b>	ASTON	87	LASS	0	11 $K^-p \rightarrow \bar{K}^0\pi^+\pi^-n$
$\Gamma(K^*(892)\pi)/\Gamma(K\pi)$					$\Gamma_2/\Gamma_3$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>1.09<math>\pm</math>0.26 OUR FIT</b>					
<b>1.09<math>\pm</math>0.26</b>	ASTON	84B	LASS	0	11 $K^-p \rightarrow \bar{K}^02\pi n$
$\Gamma(K\pi)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>0.188<math>\pm</math>0.010 OUR FIT</b>					
<b>0.188<math>\pm</math>0.010 OUR AVERAGE</b>					
0.187 $\pm$ 0.008 $\pm$ 0.008	ASTON	88	LASS	0	11 $K^-p \rightarrow K^-\pi^+n$
0.19 $\pm$ 0.02	ESTABROOKS 78	ASPK	0	13 $K^\pm p \rightarrow K\pi N$	
$\Gamma(K\eta)/\Gamma(K\pi)$					$\Gamma_4/\Gamma_3$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>1.6 <math>\pm</math> 0.7 OUR FIT</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.41 $\pm$ 0.050	<sup>12</sup> BIRD	89	LASS	—	11 $K^-p \rightarrow \bar{K}^0\pi^-p$
0.50 $\pm$ 0.18	ASTON	88B	LASS	—	11 $K^-p \rightarrow K^-\eta p$
<sup>12</sup> This result supersedes ASTON 88B.					

$\Gamma(K_2^*(1430)\pi)/\Gamma(K^*(892)\pi)$						$\Gamma_5/\Gamma_2$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<b>&lt;0.78</b>	95	ASTON	87	LASS	0	11 $K^-p \rightarrow \bar{K}^0\pi^+\pi^-n$

$K_3^*(1780)$ REFERENCES					
BIRD	89	SLAC-332	P.F. Bird		(SLAC)
ASTON	88	NP B296 493	D. Aston <i>et al.</i>		(SLAC, NAGO, CINC, INUS)
ASTON	88B	PL B201 169	D. Aston <i>et al.</i>		(SLAC, NAGO, CINC, INUS) JP
ASTON	87	NP B292 693	D. Aston <i>et al.</i>		(SLAC, NAGO, CINC, INUS)
ASTON	84B	NP B247 261	D. Aston <i>et al.</i>		(SLAC, CARL, OTTA)
BAUBILLIER	84B	ZPHY C26 37	M. Baubillier <i>et al.</i>		(BIRM, CERN, GLAS+)
BAUBILLIER	82B	NP B202 21	M. Baubillier <i>et al.</i>		(BIRM, CERN, GLAS+)
CLELAND	82	NP B208 189	W.E. Cleland <i>et al.</i>		(DURH, GEVA, LAUS+)
ASTON	81D	PL 99B 502	D. Aston <i>et al.</i>		(SLAC, CARL, OTTA) JP
TOAFF	81	PR D23 1500	S. Toaff <i>et al.</i>		(ANL, KANS)
ETKIN	80	PR D22 42	A. Etkin <i>et al.</i>		(BNL, CUNY) JP
BEUSCH	78	PL 74B 282	W. Beusch <i>et al.</i>		(CERN, AACH3, ETH) JP
CHUNG	78	PRL 40 355	S.U. Chung <i>et al.</i>		(BNL, BRAN, CUNY+) JP
ESTABROOKS	78	NP B133 490	P.G. Estabrooks <i>et al.</i>		(MCGI, CARL, DURH+) JP
Also		PR D17 658	P.G. Estabrooks <i>et al.</i>		(MCGI, CARL, DURH+)
BALDI	76	PL 63B 344	R. Baldi <i>et al.</i>		(GEVA) JP
BRANDENB...	76D	PL 60B 478	G.W. Brandenburg <i>et al.</i>		(SLAC) JP

$K_2(1820)$

$I(J^P) = \frac{1}{2}(2^-)$

See our mini-review in the 2004 edition of this *Review* (PDG 04) under  $K_2(1770)$ .

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1816<math>\pm</math>13</b>	<sup>1</sup> ASTON	93	LASS 11 $K^-p \rightarrow K^-\omega p$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$\sim 1840$	<sup>2</sup> DAUM	81c	CNTR 63 $K^-p \rightarrow K^-2\pi p$
<sup>1</sup> From a partial wave analysis of the $K^-\omega$ system.			
<sup>2</sup> From a partial wave analysis of the $K^-2\pi$ system.			

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>276<math>\pm</math>35</b>	<sup>3</sup> ASTON	93	LASS 11 $K^-p \rightarrow K^-\omega p$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$\sim 230$	<sup>4</sup> DAUM	81c	CNTR 63 $K^-p \rightarrow K^-2\pi p$
<sup>3</sup> From a partial wave analysis of the $K^-\omega$ system.			
<sup>4</sup> From a partial wave analysis of the $K^-2\pi$ system.			

$K_2(1820)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K\pi\pi$	
$\Gamma_2$ $K_2^*(1430)\pi$	seen
$\Gamma_3$ $K^*(892)\pi$	seen
$\Gamma_4$ $Kf_2(1270)$	seen
$\Gamma_5$ $K\omega$	seen

$K_2(1820)$ BRANCHING RATIOS					
$\Gamma(K_2^*(1430)\pi)/\Gamma(K\pi\pi)$					$\Gamma_2/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$\sim 0.77$	DAUM	81c	CNTR 63 $K^-p \rightarrow \bar{K}2\pi p$		
$\Gamma(K^*(892)\pi)/\Gamma(K\pi\pi)$					$\Gamma_3/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$\sim 0.05$	DAUM	81c	CNTR 63 $K^-p \rightarrow \bar{K}2\pi p$		
$\Gamma(Kf_2(1270))/\Gamma(K\pi\pi)$					$\Gamma_4/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$\sim 0.18$	DAUM	81c	CNTR 63 $K^-p \rightarrow \bar{K}2\pi p$		

$K_2(1820)$ REFERENCES					
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>		(PDG Collab.)
ASTON	93	PL B308 186	D. Aston <i>et al.</i>		(SLAC, NAGO, CINC, INUS)
DAUM	81C	NP B187 1	C. Daum <i>et al.</i>		(AMST, CERN, CRAC, MPIM+)



See key on page 601

Meson Particle Listings

$K(1830)$ ,  $K_0^*(1950)$ ,  $K_2^*(1980)$ ,  $K_4^*(2045)$

$K(1830)$

$I(J^P) = \frac{1}{2}(0^-)$

OMITTED FROM SUMMARY TABLE

Seen in partial-wave analysis of  $K^- \phi$  system. Needs confirmation.

K(1830) MASS				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
~ 1830	ARMSTRONG 83	OMEG	—	18.5 $K^- p \rightarrow 3 K p$
K(1830) WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
~ 250	ARMSTRONG 83	OMEG	—	18.5 $K^- p \rightarrow 3 K p$

K(1830) DECAY MODES	
Mode	
$\Gamma_1$ $K \phi$	

K(1830) REFERENCES			
ARMSTRONG 83	NP B221 1	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)JP

$K_0^*(1950)$

$I(J^P) = \frac{1}{2}(0^+)$

OMITTED FROM SUMMARY TABLE

Seen in partial-wave analysis of the  $K^- \pi^+$  system. Needs confirmation.

K_0^*(1950) MASS				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>1945 ± 10 ± 20</b>	<sup>1</sup> ASTON 88	LASS	0	11 $K^- p \rightarrow K^- \pi^+ n$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
1917 ± 12	<sup>2</sup> ZHOU 06	RVUE		$K p \rightarrow K^- \pi^+ n$
1820 ± 40	<sup>3</sup> ANISOVICH 97c	RVUE		11 $K^- p \rightarrow K^- \pi^+ n$
<sup>1</sup> We take the central value of the two solutions and the larger error given.				
<sup>2</sup> S-matrix pole. Using ASTON 88 and assuming $K_0^*(800)$ , $K_0^*(1430)$ .				
<sup>3</sup> T-matrix pole. Reanalysis of ASTON 88 data.				

K_0^*(1950) WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>201 ± 34 ± 79</b>	<sup>4</sup> ASTON 88	LASS	0	11 $K^- p \rightarrow K^- \pi^+ n$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
145 ± 38	<sup>5</sup> ZHOU 06	RVUE		$K p \rightarrow K^- \pi^+ n$
250 ± 100	<sup>6</sup> ANISOVICH 97c	RVUE		11 $K^- p \rightarrow K^- \pi^+ n$
<sup>4</sup> We take the central value of the two solutions and the larger error given.				
<sup>5</sup> S-matrix pole. Using ASTON 88 and assuming $K_0^*(800)$ , $K_0^*(1430)$ .				
<sup>6</sup> T-matrix pole. Reanalysis of ASTON 88 data.				

K_0^*(1950) DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K \pi$	(52 ± 14) %

K_0^*(1950) BRANCHING RATIOS				
$\Gamma(K \pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.52 ± 0.08 ± 0.12</b>	<sup>7</sup> ASTON 88	LASS	0	11 $K^- p \rightarrow K^- \pi^+ n$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
~ 0.60	<sup>8</sup> ZHOU 06	RVUE		$K p \rightarrow K^- \pi^+ n$
<sup>7</sup> We take the central value of the two solutions and the larger error given.				
<sup>8</sup> S-matrix pole. Using ASTON 88 and assuming $K_0^*(800)$ , $K_0^*(1430)$ .				

K_0^*(1950) REFERENCES			
ZHOU 06	NP A775 212	Z.Y. Zhou, H.Q. Zheng	
ANISOVICH 97c	PL B413 137	A.V. Anisovich, A.V. Sarantsev	
ASTON 88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)

$K_2^*(1980)$

$I(J^P) = \frac{1}{2}(2^+)$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

K_2^*(1980) MASS				
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG COMMENT
<b>1973 ± 8 ± 25</b>		ASTON 87	LASS	0 11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
2020 ± 20		TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
1978 ± 40	241 ± 47	BIRD 89	LASS	— 11 $K^- p \rightarrow \bar{K}^0 \pi^- p$

K_2^*(1980) WIDTH				
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG COMMENT
<b>373 ± 33 ± 60</b>		ASTON 87	LASS	0 11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
180 ± 70		TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
398 ± 47	241 ± 47	BIRD 89	LASS	— 11 $K^- p \rightarrow \bar{K}^0 \pi^- p$

K_2^*(1980) DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $K^*(892) \pi$	possibly seen
$\Gamma_2$ $K \rho$	possibly seen
$\Gamma_3$ $K f_2(1270)$	possibly seen

K_2^*(1980) BRANCHING RATIOS				
$\Gamma(K^*(892) \pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
possibly seen	GULER 11	BELL	$B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$	
$\Gamma(K \rho)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
possibly seen	GULER 11	BELL	$B^+ \rightarrow J/\psi K^+ \pi^+ \pi^-$	
$\Gamma(K \rho)/\Gamma(K^*(892) \pi)$				$\Gamma_2/\Gamma_1$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.49 ± 0.24 ± 0.09</b>	ASTON 87	LASS	0	11 $K^- p \rightarrow \bar{K}^0 \pi^+ \pi^- n$
$\Gamma(K f_2(1270))/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
possibly seen	TIKHOMIROV 03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$	

K_2^*(1980) REFERENCES			
GULER 11	PR D83 032005	H. Guler <i>et al.</i>	(BELLE Collab.)
TIKHOMIROV 03	PAN 66 928	G.D. Tikhomirov <i>et al.</i>	
BIRD 89	SLAC-332	P.F. Bird	(SLAC)
ASTON 87	NP B292 693	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)

$K_4^*(2045)$

$I(J^P) = \frac{1}{2}(4^+)$

K_4^*(2045) MASS				
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG COMMENT
<b>2045 ± 9 OUR AVERAGE</b>		Error includes scale factor of 1.1.		
2062 ± 14 ± 13		<sup>1</sup> ASTON 86	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$
2039 ± 10	400	<sup>2,3</sup> CLELAND 82	SPEC	± 50 $K^+ p \rightarrow K_S^0 \pi^\pm p$
2070 <sup>+100</sup> <sub>-40</sub>		<sup>4</sup> ASTON 81c	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
2079 ± 7	431	TORRES 86	MPSF	400 $p A \rightarrow 4 K X$
2088 ± 20	650	BAUBILLIER 82	HBC	— 8.25 $K^- p \rightarrow K_S^0 \pi^- p$
2115 ± 46	488	CARMONY 77	HBC	0 9 $K^+ d \rightarrow K^+ \pi^+ X$
<sup>1</sup> From a fit to all moments.				
<sup>2</sup> From a fit to 8 moments.				
<sup>3</sup> Number of events evaluated by us.				
<sup>4</sup> From energy-independent partial-wave analysis.				

K_4^*(2045) WIDTH				
VALUE (MeV)	EVTs	DOCUMENT ID	TECN	CHG COMMENT
<b>198 ± 30 OUR AVERAGE</b>				
221 ± 48 ± 27		<sup>5</sup> ASTON 86	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$
189 ± 35	400	<sup>6,7</sup> CLELAND 82	SPEC	± 50 $K^+ p \rightarrow K_S^0 \pi^\pm p$

Meson Particle Listings

$K_4^*(2045)$ ,  $K_2(2250)$ ,  $K_3(2320)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$61 \pm 58$	431	TORRES	86	MPSF	400 $pA \rightarrow 4 K X$
$170^{+100}_{-50}$	650	BAUBILLIER	82	HBC	— 8.25 $K^- p \rightarrow K_S^0 \pi^- p$
$240^{+500}_{-100}$		<sup>8</sup> ASTON	81c	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$
300±200		CARMONY	77	HBC	0 9 $K^+ d \rightarrow K^+ \pi^- s X$

<sup>5</sup> From a fit to all moments.  
<sup>6</sup> From a fit to 8 moments.  
<sup>7</sup> Number of events evaluated by us.  
<sup>8</sup> From energy-independent partial-wave analysis.

$K_4^*(2045)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $K \pi$	(9.9±1.2) %	
$\Gamma_2$ $K^*(892) \pi \pi$	(9 ±5 ) %	
$\Gamma_3$ $K^*(892) \pi \pi \pi$	(7 ±5 ) %	
$\Gamma_4$ $\rho K \pi$	(5.7±3.2) %	
$\Gamma_5$ $\omega K \pi$	(5.0±3.0) %	
$\Gamma_6$ $\phi K \pi$	(2.8±1.4) %	
$\Gamma_7$ $\phi K^*(892)$	(1.4±0.7) %	

$K_4^*(2045)$ BRANCHING RATIOS				
$\Gamma(K \pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.099±0.012</b>	ASTON	88	LASS	0 11 $K^- p \rightarrow K^- \pi^+ n$
$\Gamma(K^*(892) \pi \pi)/\Gamma(K \pi)$	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.89±0.53</b>	BAUBILLIER	82	HBC	— 8.25 $K^- p \rightarrow p K_S^0 3 \pi$
$\Gamma(K^*(892) \pi \pi \pi)/\Gamma(K \pi)$	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.75±0.49</b>	BAUBILLIER	82	HBC	— 8.25 $K^- p \rightarrow p K_S^0 3 \pi$
$\Gamma(\rho K \pi)/\Gamma(K \pi)$	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.58±0.32</b>	BAUBILLIER	82	HBC	— 8.25 $K^- p \rightarrow p K_S^0 3 \pi$
$\Gamma(\omega K \pi)/\Gamma(K \pi)$	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.50±0.30</b>	BAUBILLIER	82	HBC	— 8.25 $K^- p \rightarrow p K_S^0 3 \pi$
$\Gamma(\phi K \pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
<b>0.028±0.014</b>	<sup>9</sup> TORRES	86	MPSF	400 $pA \rightarrow 4 K X$
$\Gamma(\phi K^*(892))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
<b>0.014±0.007</b>	<sup>9</sup> TORRES	86	MPSF	400 $pA \rightarrow 4 K X$

<sup>9</sup> Error determination is model dependent.

$K_4^*(2045)$ REFERENCES				
ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	86	PL B180 308	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
TORRES	86	PR D34 707	S. Torres <i>et al.</i>	(VPI, ARIZ, FNAL, FSU+)
BAUBILLIER	82	PL 118B 447	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
CLELAND	82	NP B208 189	W.E. Cleland <i>et al.</i>	(DURH, GEVA, LAUS+)
ASTON	81c	PL 106B 235	D. Aston <i>et al.</i>	(SLAC, CARL, OTTA) JP
CARMONY	77	PR D16 1251	D.D. Carmony <i>et al.</i>	(PURD, UCD, IUPU)

$K_2(2250)$	$I(J^P) = \frac{1}{2}(2^-)$
OMITTED FROM SUMMARY TABLE	
This entry contains various peaks in strange meson systems reported in the 2150–2260 MeV region, as well as enhancements seen in the antihyperon-nucleon system, either in the mass spectra or in the $J^P = 2^-$ wave.	

$K_2(2250)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG COMMENT
<b>2247±17 OUR AVERAGE</b>				
2200±40		<sup>1</sup> ARMSTRONG	83c	OMEG — 18 $K^- p \rightarrow \Lambda \bar{p} X$
2235±50		<sup>1</sup> BAUBILLIER	81	HBC — 8 $K^- p \rightarrow \Lambda \bar{p} X$
2260±20		<sup>1</sup> CLELAND	81	SPEC ± 50 $K^+ p \rightarrow \Lambda \bar{p} X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

2280±20		TIKHOMIROV	03	SPEC	40.0 $\pi^- C \rightarrow K_S^0 K_S^0 K_L^0 X$
2147± 4	37	CHLIAPNIK...	79	HBC	+ 32 $K^+ p \rightarrow \Lambda \bar{p} X$
2240±20	20	LISSAUER	70	HBC	9 $K^+ p$

<sup>1</sup>  $J^P = 2^-$  from moments analysis.

$K_2(2250)$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG COMMENT
<b>180±30 OUR AVERAGE</b>				Error includes scale factor of 1.4.
150±30		<sup>2</sup> ARMSTRONG	83c	OMEG — 18 $K^- p \rightarrow \Lambda \bar{p} X$
210±30		<sup>2</sup> CLELAND	81	SPEC ± 50 $K^+ p \rightarrow \Lambda \bar{p} X$
180±60		TIKHOMIROV	03	SPEC
~ 200		<sup>2</sup> BAUBILLIER	81	HBC — 8 $K^- p \rightarrow \Lambda \bar{p} X$
~ 40	37	CHLIAPNIK...	79	HBC + 32 $K^+ p \rightarrow \Lambda \bar{p} X$
80±20	20	LISSAUER	70	HBC
				9 $K^+ p$

<sup>2</sup>  $J^P = 2^-$  from moments analysis.

$K_2(2250)$ DECAY MODES				
Mode				
$\Gamma_1$ $K \pi \pi$				
$\Gamma_2$ $K f_2(1270)$				
$\Gamma_3$ $K^*(892) f_0(980)$				
$\Gamma_4$ $\rho \bar{\Lambda}$				

$K_2(2250)$ REFERENCES				
TIKHOMIROV	03	PAN 66 828	G.D. Tikhomirov <i>et al.</i>	
ARMSTRONG	83c	Translated from YAF 66 860.	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
BAUBILLIER	81	NP B183 1	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
CLELAND	81	NP B184 1	W.E. Cleland <i>et al.</i>	(PITT, GEVA, LAUS+)
CHLIAPNIK...	79	NP B158 253	P.V. Chliapnikov <i>et al.</i>	(CERN, BELG, MONS)
LISSAUER	70	NP B18 491	D. Lissauer <i>et al.</i>	(LBL)

$K_3(2320)$	$I(J^P) = \frac{1}{2}(3^+)$
OMITTED FROM SUMMARY TABLE	
Seen in the $J^P = 3^+$ wave of the antihyperon-nucleon system.	
Needs confirmation.	

$K_3(2320)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>2324±24 OUR AVERAGE</b>				
2330±40		<sup>1</sup> ARMSTRONG	83c	OMEG — 18 $K^- p \rightarrow \Lambda \bar{p} X$
2320±30		<sup>1</sup> CLELAND	81	SPEC ± 50 $K^+ p \rightarrow \Lambda \bar{p} X$

<sup>1</sup>  $J^P = 3^+$  from moments analysis.

$K_3(2320)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>150±30</b>		<sup>2</sup> ARMSTRONG	83c	OMEG — 18 $K^- p \rightarrow \Lambda \bar{p} X$
~ 250		<sup>2</sup> CLELAND	81	SPEC ± 50 $K^+ p \rightarrow \Lambda \bar{p} X$

<sup>2</sup>  $J^P = 3^+$  from moments analysis.

$K_3(2320)$ DECAY MODES				
Mode				
$\Gamma_1$	$\rho \overline{\Lambda}$			
$K_3(2320)$ REFERENCES				
ARMSTRONG	83c	NP B227 365	T.A. Armstrong <i>et al.</i>	(BARI, BIRM, CERN+)
CLELAND	81	NP B184 1	W.E. Cleland <i>et al.</i>	(PITT, GEVA, LAUS+)

See key on page 601

Meson Particle Listings  
 $K_5^*(2380)$ ,  $K_4(2500)$ ,  $K(3100)$

$K_5^*(2380)$

$I(J^P) = \frac{1}{2}(5^-)$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

$K_5^*(2380)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>2382±14±19</b>	<sup>1</sup> ASTON	86	LASS	0	11 $K^-p \rightarrow K^- \pi^+ n$

<sup>1</sup>From a fit to all the moments.

$K_5^*(2380)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT	
<b>178±37±32</b>	<sup>2</sup> ASTON	86	LASS	0	11 $K^-p \rightarrow K^- \pi^+ n$

<sup>2</sup>From a fit to all the moments.

$K_5^*(2380)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad K \pi$	(6.1 $\pm$ 1.2) %

$K_5^*(2380)$  BRANCHING RATIOS

$\Gamma(K \pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_1/\Gamma$
$0.061 \pm 0.012$	ASTON	88	LASS	0	11 $K^- p \rightarrow K^- \pi^+ n$

$K_5^*(2380)$  REFERENCES

ASTON	88	NP B296 493	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
ASTON	86	PL B180 308	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)

$K_4(2500)$

$I(J^P) = \frac{1}{2}(4^-)$

OMITTED FROM SUMMARY TABLE

Needs confirmation.

$K_4(2500)$  MASS

VALUE (MeV)	DOCUMENT ID		TECN	CHG	COMMENT
<b>2490 ± 20</b>	<sup>1</sup> CLELAND	81	SPEC	±	50 $K^+ p \rightarrow \Lambda \bar{p}$

<sup>1</sup>  $J^P = 4^-$  from moments analysis.

$K_4(2500)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
~ 250	<sup>2</sup> CLELAND	81	SPEC	± 50 K <sup>+</sup> p → Λ p̄
<sup>2</sup> J <sup>P</sup> = 4 <sup>-</sup> from moments analysis.				

$K_4(2500)$  DECAY MODES

Mode
$\Gamma_1 \quad \rho \bar{\Lambda}$

$K_4(2500)$  REFERENCES

CLELAND	81	NP B184 1	W.E. Cleland <i>et al.</i>	(PITT, GEVA, LAUS+)
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$K(3100)$

$I^G(J^{PC}) = ?^?(?^{??})$

OMITTED FROM SUMMARY TABLE

Narrow peak observed in several ( $\Lambda \bar{p}$  + pions) and ( $\bar{\Lambda} p$  + pions) states in  $\Sigma^-$  Be reactions by BOURQUIN 86 and in  $n p$  and  $n A$  reactions by ALEEV 93. Not seen by BOEHNLEIN 91. If due to strong decays, this state has exotic quantum numbers ( $B=0, Q=+1, S=-1$  for  $\Lambda \bar{p} \pi^+ \pi^+$  and  $I \geq 3/2$  for  $\Lambda \bar{p} \pi^-$ ). Needs confirmation.

$K(3100)$  MASS

VALUE (MeV)	DOCUMENT ID
$\approx 3100$ OUR ESTIMATE	

3-BODY DECAYS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>3054 ± 11 OUR AVERAGE</b>				
3060 ± 7 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^+$
3056 ± 7 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \bar{\Lambda} p \pi^-$
3055 ± 8 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^-$
3045 ± 8 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \bar{\Lambda} p \pi^+$

4-BODY DECAYS

VALUE [MeV]	DOCUMENT ID		TECN	COMMENT
<b>3059 ± 11 OUR AVERAGE</b>				
3067 ± 6 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^+$
3060 ± 8 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^-$
3055 ± 7 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^- \pi^-$
3052 ± 8 ± 20	<sup>1</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^- \pi^+$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
3105 ± 30	BOURQUIN	86	SPEC	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^+$
3115 ± 30	BOURQUIN	86	SPEC	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^-$

5-BODY DECAYS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
3095 ± 30	BOURQUIN	86	SPEC	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^+ \pi^-$
<sup>1</sup> Supersedes ALEEV 90.				

$K(3100)$  WIDTH

3-BODY DECAYS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
42 ± 16	<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^+$
36 ± 15	<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \bar{\Lambda} p \pi^-$
50 ± 18	<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^-$
30 ± 15	<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \bar{\Lambda} p \pi^+$

4-BODY DECAYS

VALUE [MeV]	CL%	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
22 ± 8		<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^+$
28 ± 12		<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^-$
32 ± 15		<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \bar{\Lambda} p \pi^- \pi^-$
30 ± 15		<sup>2</sup> ALEEV	93	BIS2	$K(3100) \rightarrow \bar{\Lambda} p \pi^- \pi^+$
<30	90	BOURQUIN	86	SPEC	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^+$
<80	90	BOURQUIN	86	SPEC	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^-$

5-BODY DECAYS

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
<30	90	BOURQUIN	86	SPEC	$K(3100) \rightarrow \Lambda \bar{p} \pi^+ \pi^+ \pi^-$
<sup>2</sup> Supersedes ALEEV 90.					

$K(3100)$  DECAY MODES

Mode
$\Gamma_1 \quad K(3100)^0 \rightarrow \Lambda \bar{p} \pi^+$
$\Gamma_2 \quad K(3100)^-- \rightarrow \Lambda \bar{p} \pi^-$
$\Gamma_3 \quad K(3100)^- \rightarrow \Lambda \bar{p} \pi^+ \pi^-$
$\Gamma_4 \quad K(3100)^+ \rightarrow \Lambda \bar{p} \pi^+ \pi^+$
$\Gamma_5 \quad K(3100)^0 \rightarrow \Lambda \bar{p} \pi^+ \pi^+ \pi^-$
$\Gamma_6 \quad K(3100)^0 \rightarrow \Sigma(1385)^+ \bar{p}$

$\Gamma(\Sigma(1385)^+\bar{p})/\Gamma(\Lambda\bar{p}\pi^+)$					$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.04	90	ALEEV	93	BIS2	$K(3100)^0 \rightarrow \Sigma(1385)^+ \bar{p}$

$K(3100)$  REFERENCES

ALEEV	93	PAN 56 1358	A.N. Aleev <i>et al.</i>	(BIS-2 Collab.)
BOEHNLEIN	91	NPBPS B21 174	A. Boehnlein <i>et al.</i>	(FLOR, BNL, IND+)
ALEEV	90	ZPHY C47 533	A.N. Aleev <i>et al.</i>	(BIS-2 Collab.)
BOURQUIN	86	PL B172 113	M.H. Bourquin <i>et al.</i>	(GEVA, RAL, HEIDP+)

# Meson Particle Listings

$D^\pm$

CHARMED MESONS  
( $C = \pm 1$ )  
 $D^+ = c\bar{d}$ ,  $D^0 = c\bar{u}$ ,  $\bar{D}^0 = \bar{c}u$ ,  $D^- = \bar{c}d$ , similarly for  $D^{*}$ 's

$D^\pm$

$I(J^P) = \frac{1}{2}(0^-)$

$D^\pm$  MASS

The fit includes  $D^\pm$ ,  $D^0$ ,  $D_S^\pm$ ,  $D^{*\pm}$ ,  $D^{*0}$ ,  $D_S^{*\pm}$ ,  $D_1(2420)^0$ ,  $D_2^*(2460)^0$ , and  $D_{S1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1869.58 ± 0.09 OUR FIT</b>				
<b>1869.5 ± 0.4 OUR AVERAGE</b>				
1869.53 ± 0.49 ± 0.20	110 ± 15	ANASHIN	10A	KEDR $e^+e^-$ at $\psi(3770)$
1870.0 ± 0.5 ± 1.0	317	BARLAG	90c	ACCM $\pi^-$ Cu 230 GeV
1869.4 ± 0.6		<sup>1</sup> TRILLING	81	RVUE $e^+e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1875 ± 10	9	ADAMOVIICH	87	EMUL Photoproduction
1860 ± 16	6	ADAMOVIICH	84	EMUL Photoproduction
1863 ± 4		DERRICK	84	HRS $e^+e^-$ 29 GeV
1868.4 ± 0.5		<sup>1</sup> SCHINDLER	81	MRK2 $e^+e^-$ 3.77 GeV
1874 ± 5		GOLDHABER	77	MRK1 $D^0$ , $D^+$ recoil spectra
1868.3 ± 0.9		<sup>1</sup> PERUZZI	77	LGW $e^+e^-$ 3.77 GeV
1874 ± 11		PICCOLO	77	MRK1 $e^+e^-$ 4.03, 4.41 GeV
1876 ± 15	50	PERUZZI	76	MRK1 $K^\mp\pi^\pm\pi^\pm$

<sup>1</sup> PERUZZI 77 and SCHINDLER 81 errors do not include the 0.13% uncertainty in the absolute SPEAR energy calibration. TRILLING 81 uses the high precision  $J/\psi(1S)$  and  $\psi(2S)$  measurements of ZHOLENTZ 80 to determine this uncertainty and combines the PERUZZI 77 and SCHINDLER 81 results to obtain the value quoted.

$D^\pm$  MEAN LIFE

Measurements with an error  $> 100 \times 10^{-15}$  s have been omitted from the Listings.

VALUE ( $10^{-15}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1040 ± 7 OUR AVERAGE</b>				
1039.4 ± 4.3 ± 7.0	110k	LINK	02F	FOCS $\gamma$ nucleus, $\approx 180$ GeV
1033.6 ± 22.1 ± 9.9 12.7	3777	BONVICINI	99	CLEO $e^+e^- \approx \Upsilon(4S)$
1048 ± 15 ± 11	9k	FRABETTI	94d	E687 $D^+ \rightarrow K^- \pi^+ \pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1075 ± 40 ± 18	2455	FRABETTI	91	E687 $\gamma$ Be, $D^+ \rightarrow K^- \pi^+ \pi^+$
1030 ± 80 ± 60	200	ALVAREZ	90	NA14 $\gamma$ , $D^+ \rightarrow K^- \pi^+ \pi^+$
1050 ± 77 -72	317	<sup>1</sup> BARLAG	90c	ACCM $\pi^-$ Cu 230 GeV
1050 ± 80 ± 70	363	ALBRECHT	88i	ARG $e^+e^-$ 10 GeV
1090 ± 30 ± 25	2992	RAAB	88	E691 Photoproduction

<sup>1</sup> BARLAG 90c estimates the systematic error to be negligible.

$D^+$  DECAY MODES

Most decay modes (other than the semileptonic modes) that involve a neutral  $K$  meson are now given as  $K_S^0$  modes, not as  $\bar{K}^0$  modes. Nearly always it is a  $K_S^0$  that is measured, and interference between Cabibbo-allowed and doubly Cabibbo-suppressed modes can invalidate the assumption that  $2\Gamma(K_S^0) = \Gamma(\bar{K}^0)$ .

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Inclusive modes</b>		
$\Gamma_1$ $e^+$ semileptonic	(16.07 ± 0.30) %	
$\Gamma_2$ $\mu^+$ anything	(17.6 ± 3.2) %	
$\Gamma_3$ $K^-$ anything	(25.7 ± 1.4) %	
$\Gamma_4$ $\bar{K}^0$ anything + $K^0$ anything	(61 ± 5) %	
$\Gamma_5$ $K^+$ anything	( 5.9 ± 0.8) %	
$\Gamma_6$ $K^*(892)^-$ anything	( 6 ± 5) %	
$\Gamma_7$ $\bar{K}^*(892)^0$ anything	(23 ± 5) %	
$\Gamma_8$ $K^*(892)^0$ anything	< 6.6 %	CL=90%
$\Gamma_9$ $\eta$ anything	( 6.3 ± 0.7) %	
$\Gamma_{10}$ $\eta'$ anything	( 1.04 ± 0.18) %	
$\Gamma_{11}$ $\phi$ anything	( 1.03 ± 0.12) %	

Leptonic and semileptonic modes

$\Gamma_{12}$ $e^+ \nu_e$	< 8.8	$\times 10^{-6}$	CL=90%
$\Gamma_{13}$ $\mu^+ \nu_\mu$	( 3.74 ± 0.17)	$\times 10^{-4}$	
$\Gamma_{14}$ $\tau^+ \nu_\tau$	< 1.2	$\times 10^{-3}$	CL=90%
$\Gamma_{15}$ $\bar{K}^0 e^+ \nu_e$	( 8.90 ± 0.15) %		
$\Gamma_{16}$ $\bar{K}^0 \mu^+ \nu_\mu$	( 9.3 ± 0.7) %		
$\Gamma_{17}$ $K^- \pi^+ e^+ \nu_e$	( 3.91 ± 0.11) %		
$\Gamma_{18}$ $\bar{K}^*(892)^0 e^+ \nu_e$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 3.68 ± 0.10) %		
$\Gamma_{19}$ $(K^- \pi^+)_{S\text{-wave}} e^+ \nu_e$	( 2.26 ± 0.11)	$\times 10^{-3}$	
$\Gamma_{20}$ $\bar{K}^*(1410)^0 e^+ \nu_e$ , $\bar{K}^*(1410)^0 \rightarrow K^- \pi^+$	< 6	$\times 10^{-3}$	CL=90%
$\Gamma_{21}$ $\bar{K}_2^*(1430)^0 e^+ \nu_e$ , $\bar{K}_2^*(1430)^0 \rightarrow K^- \pi^+$	< 5	$\times 10^{-4}$	CL=90%
$\Gamma_{22}$ $K^- \pi^+ e^+ \nu_e$ nonresonant	< 7	$\times 10^{-3}$	CL=90%
$\Gamma_{23}$ $K^- \pi^+ \mu^+ \nu_\mu$	( 3.9 ± 0.4) %		
$\Gamma_{24}$ $\bar{K}^*(892)^0 \mu^+ \nu_\mu$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 3.52 ± 0.10) %		
$\Gamma_{25}$ $K^- \pi^+ \mu^+ \nu_\mu$ nonresonant	( 2.1 ± 0.5)	$\times 10^{-3}$	
$\Gamma_{26}$ $K^- \pi^+ \pi^0 \mu^+ \nu_\mu$	< 1.6	$\times 10^{-3}$	CL=90%
$\Gamma_{27}$ $\pi^0 e^+ \nu_e$	( 4.05 ± 0.18)	$\times 10^{-3}$	
$\Gamma_{28}$ $\eta e^+ \nu_e$	( 1.14 ± 0.10)	$\times 10^{-3}$	
$\Gamma_{29}$ $\rho^0 e^+ \nu_e$	( 2.18 ± 0.17 -0.25)	$\times 10^{-3}$	
$\Gamma_{30}$ $\rho^0 \mu^+ \nu_\mu$	( 2.4 ± 0.4)	$\times 10^{-3}$	
$\Gamma_{31}$ $\omega e^+ \nu_e$	( 1.69 ± 0.11)	$\times 10^{-3}$	
$\Gamma_{32}$ $\eta'(958) e^+ \nu_e$	( 2.2 ± 0.5)	$\times 10^{-4}$	
$\Gamma_{33}$ $\phi e^+ \nu_e$	< 1.3	$\times 10^{-5}$	CL=90%

Fractions of some of the following modes with resonances have already appeared above as submodes of particular charged-particle modes.

$\Gamma_{34}$ $\bar{K}^*(892)^0 e^+ \nu_e$	( 5.52 ± 0.15) %		
$\Gamma_{35}$ $\bar{K}^*(892)^0 \mu^+ \nu_\mu$	( 5.30 ± 0.15) %		
$\Gamma_{36}$ $\bar{K}_0^*(1430)^0 \mu^+ \nu_\mu$	< 2.5	$\times 10^{-4}$	CL=90%
$\Gamma_{37}$ $\bar{K}^*(1680)^0 \mu^+ \nu_\mu$	< 1.6	$\times 10^{-3}$	CL=90%

Hadronic modes with a  $\bar{K}$  or  $\bar{K}K\bar{K}$

$\Gamma_{38}$ $K^0 \pi^+$	( 1.53 ± 0.06) %		S=2.8
$\Gamma_{39}$ $K_L^0 \pi^+$	( 1.46 ± 0.05) %		
$\Gamma_{40}$ $K^- 2\pi^+$	[a] ( 9.46 ± 0.24) %		S=2.0
$\Gamma_{41}$ $(K^- \pi^+)_{S\text{-wave}} \pi^+$	( 7.58 ± 0.22) %		
$\Gamma_{42}$ $\bar{K}_0^*(800)^0 \pi^+$ , $\bar{K}_0^*(800)^0 \rightarrow K^- \pi^+$			
$\Gamma_{43}$ $\bar{K}_0^*(1430)^0 \pi^+$ , $\bar{K}_0^*(1430)^0 \rightarrow K^- \pi^+$	[b] ( 1.26 ± 0.07) %		
$\Gamma_{44}$ $\bar{K}^*(892)^0 \pi^+$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.05 ± 0.12) %		
$\Gamma_{45}$ $\bar{K}^*(1410)^0 \pi^+$ , $\bar{K}^{*0} \rightarrow K^- \pi^+$	not seen		
$\Gamma_{46}$ $\bar{K}_2^*(1430)^0 \pi^+$ , $\bar{K}_2^*(1430)^0 \rightarrow K^- \pi^+$	[b] ( 2.3 ± 0.8)	$\times 10^{-4}$	
$\Gamma_{47}$ $\bar{K}^*(1680)^0 \pi^+$ , $\bar{K}^*(1680)^0 \rightarrow K^- \pi^+$	[b] ( 2.2 ± 1.1)	$\times 10^{-4}$	
$\Gamma_{48}$ $K^-(2\pi^+)_{I=2}$	( 1.47 ± 0.27) %		
$\Gamma_{49}$ $K^- 2\pi^+$ nonresonant			
$\Gamma_{50}$ $K_S^0 \pi^+ \pi^0$	[a] ( 7.24 ± 0.17) %		
$\Gamma_{51}$ $K_S^0 \rho^+$	( 6.04 ± 0.60 -0.34) %		
$\Gamma_{52}$ $K_S^0 \rho(1450)^+$ , $\rho^+ \rightarrow \pi^+ \pi^0$	( 1.5 ± 1.2 -1.4)	$\times 10^{-3}$	
$\Gamma_{53}$ $\bar{K}^*(892)^0 \pi^+$ , $\bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0$	( 2.59 ± 0.31)	$\times 10^{-3}$	
$\Gamma_{54}$ $\bar{K}_0^*(1430)^0 \pi^+$ , $\bar{K}_0^{*0} \rightarrow K_S^0 \pi^0$	( 2.7 ± 0.9)	$\times 10^{-3}$	
$\Gamma_{55}$ $\bar{K}_0^*(1680)^0 \pi^+$ , $\bar{K}_0^{*0} \rightarrow K_S^0 \pi^0$	( 9 ± 7 -10)	$\times 10^{-4}$	
$\Gamma_{56}$ $\bar{\kappa}^0 \pi^+$ , $\bar{\kappa}^0 \rightarrow K_S^0 \pi^0$	( 6 ± 5 -4)	$\times 10^{-3}$	
$\Gamma_{57}$ $K_S^0 \pi^+ \pi^0$ nonresonant	( 3 ± 4)	$\times 10^{-3}$	
$\Gamma_{58}$ $K_S^0 \pi^+ \pi^0$ nonresonant and $\bar{\kappa}^0 \pi^+$	( 1.35 ± 0.21 -0.40) %		
$\Gamma_{59}$ $(K_S^0 \pi^0)_{S\text{-wave}} \pi^+$	( 1.25 ± 0.27 -0.33) %		
$\Gamma_{60}$ $K^- 2\pi^+ \pi^0$	[c] ( 6.14 ± 0.16) %		
$\Gamma_{61}$ $K_S^0 2\pi^+ \pi^-$	[c] ( 3.05 ± 0.09) %		

See key on page 601

## Meson Particle Listings

 $D^\pm$ 

$\Gamma_{62}$	$K^- 3\pi^+ \pi^-$	[a]	$(5.8 \pm 0.5) \times 10^{-3}$	S=1.1
$\Gamma_{63}$	$\bar{K}^*(892)^0 2\pi^+ \pi^-$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$		$(1.2 \pm 0.4) \times 10^{-3}$	
$\Gamma_{64}$	$\bar{K}^*(892)^0 \rho^0 \pi^+$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$		$(2.3 \pm 0.4) \times 10^{-3}$	
$\Gamma_{65}$	$\bar{K}^*(892)^0 a_1(1260)^+$	[d]	$(9.4 \pm 1.9) \times 10^{-3}$	
$\Gamma_{66}$	$\bar{K}^*(892)^0 2\pi^+ \pi^-$ non- $\rho$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$			
$\Gamma_{67}$	$K^- \rho^0 2\pi^+$		$(1.74 \pm 0.28) \times 10^{-3}$	
$\Gamma_{68}$	$K^- 3\pi^+ \pi^-$ nonresonant		$(4.1 \pm 3.0) \times 10^{-4}$	
$\Gamma_{69}$	$K^+ 2K_S^0$		$(4.6 \pm 2.1) \times 10^{-3}$	
$\Gamma_{70}$	$K^+ K^- K_S^0 \pi^+$		$(2.3 \pm 0.5) \times 10^{-4}$	

## Pionic modes

$\Gamma_{71}$	$\pi^+ \pi^0$		$(1.24 \pm 0.06) \times 10^{-3}$	
$\Gamma_{72}$	$2\pi^+ \pi^-$		$(3.29 \pm 0.20) \times 10^{-3}$	
$\Gamma_{73}$	$\rho^0 \pi^+$		$(8.4 \pm 1.5) \times 10^{-4}$	
$\Gamma_{74}$	$\pi^+ (\pi^+ \pi^-)_{S\text{-wave}}$		$(1.85 \pm 0.17) \times 10^{-3}$	
$\Gamma_{75}$	$\sigma \pi^+, \sigma \rightarrow \pi^+ \pi^-$		$(1.39 \pm 0.12) \times 10^{-3}$	
$\Gamma_{76}$	$f_0(980) \pi^+$ , $f_0(980) \rightarrow \pi^+ \pi^-$		$(1.58 \pm 0.34) \times 10^{-4}$	
$\Gamma_{77}$	$f_0(1370) \pi^+$ , $f_0(1370) \rightarrow \pi^+ \pi^-$		$(8 \pm 4) \times 10^{-5}$	
$\Gamma_{78}$	$f_2(1270) \pi^+$ , $f_2(1270) \rightarrow \pi^+ \pi^-$		$(5.1 \pm 0.9) \times 10^{-4}$	
$\Gamma_{79}$	$\rho(1450)^0 \pi^+$ , $\rho(1450)^0 \rightarrow \pi^+ \pi^-$	$< 8$	$\times 10^{-5}$	CL=95%
$\Gamma_{80}$	$f_0(1500) \pi^+$ , $f_0(1500) \rightarrow \pi^+ \pi^-$		$(1.1 \pm 0.4) \times 10^{-4}$	
$\Gamma_{81}$	$f_0(1710) \pi^+$ , $f_0(1710) \rightarrow \pi^+ \pi^-$	$< 5$	$\times 10^{-5}$	CL=95%
$\Gamma_{82}$	$f_0(1790) \pi^+$ , $f_0(1790) \rightarrow \pi^+ \pi^-$	$< 7$	$\times 10^{-5}$	CL=95%
$\Gamma_{83}$	$(\pi^+ \pi^+)_{S\text{-wave}} \pi^-$	$< 1.2$	$\times 10^{-4}$	CL=95%
$\Gamma_{84}$	$2\pi^+ \pi^-$ nonresonant	$< 1.2$	$\times 10^{-4}$	CL=95%
$\Gamma_{85}$	$\pi^+ 2\pi^0$		$(4.7 \pm 0.4) \times 10^{-3}$	
$\Gamma_{86}$	$2\pi^+ \pi^- \pi^0$		$(1.17 \pm 0.08) \%$	
$\Gamma_{87}$	$\eta \pi^+, \eta \rightarrow \pi^+ \pi^- \pi^0$		$(8.0 \pm 0.5) \times 10^{-4}$	
$\Gamma_{88}$	$\omega \pi^+, \omega \rightarrow \pi^+ \pi^- \pi^0$	$< 3$	$\times 10^{-4}$	CL=90%
$\Gamma_{89}$	$3\pi^+ 2\pi^-$		$(1.67 \pm 0.16) \times 10^{-3}$	

Fractions of some of the following modes with resonances have already appeared above as submodes of particular charged-particle modes.

$\Gamma_{90}$	$\eta \pi^+$		$(3.66 \pm 0.22) \times 10^{-3}$	
$\Gamma_{91}$	$\eta \pi^+ \pi^0$		$(1.38 \pm 0.35) \times 10^{-3}$	
$\Gamma_{92}$	$\omega \pi^+$	$< 3.4$	$\times 10^{-4}$	CL=90%
$\Gamma_{93}$	$\eta'(958) \pi^+$		$(4.84 \pm 0.31) \times 10^{-3}$	
$\Gamma_{94}$	$\eta'(958) \pi^+ \pi^0$		$(1.6 \pm 0.5) \times 10^{-3}$	

Hadronic modes with a  $K\bar{K}$  pair

$\Gamma_{95}$	$K^+ K_S^0$		$(2.95 \pm 0.15) \times 10^{-3}$	S=2.8
$\Gamma_{96}$	$K^+ K^- \pi^+$	[a]	$(9.96 \pm 0.26) \times 10^{-3}$	S=1.3
$\Gamma_{97}$	$\phi \pi^+, \phi \rightarrow K^+ K^-$		$(2.77 \pm_{-0.10}^{+0.09}) \times 10^{-3}$	
$\Gamma_{98}$	$K^+ \bar{K}^*(892)^0$ , $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$		$(2.56 \pm_{-0.15}^{+0.09}) \times 10^{-3}$	
$\Gamma_{99}$	$K^+ \bar{K}_0^*(1430)^0, \bar{K}_0^*(1430)^0 \rightarrow$ $K^- \pi^+$		$(1.9 \pm 0.4) \times 10^{-3}$	
$\Gamma_{100}$	$K^+ \bar{K}_2^*(1430)^0, \bar{K}_2^* \rightarrow$ $K^- \pi^+$		$(1.7 \pm_{-0.8}^{+1.3}) \times 10^{-4}$	
$\Gamma_{101}$	$K^+ \bar{K}_0^*(800), \bar{K}_0^* \rightarrow K^- \pi^+$		$(7.0 \pm_{-2.2}^{+4.0}) \times 10^{-4}$	
$\Gamma_{102}$	$a_0(1450)^0 \pi^+, a_0^0 \rightarrow K^+ K^-$		$(4.6 \pm_{-1.9}^{+7.0}) \times 10^{-4}$	
$\Gamma_{103}$	$\phi(1680) \pi^+, \phi \rightarrow K^+ K^-$		$(5.1 \pm_{-1.9}^{+4.0}) \times 10^{-5}$	
$\Gamma_{104}$	$K^+ K^- \pi^+$ nonresonant	not seen		
$\Gamma_{105}$	$K^+ K_S^0 \pi^+ \pi^-$		$(1.71 \pm 0.18) \times 10^{-3}$	
$\Gamma_{106}$	$K_S^0 K^- 2\pi^+$		$(2.34 \pm 0.17) \times 10^{-3}$	
$\Gamma_{107}$	$K^+ K^- 2\pi^+ \pi^-$		$(2.3 \pm 1.2) \times 10^{-4}$	

A few poorly measured branching fractions:

$\Gamma_{108}$	$\phi \pi^+ \pi^0$		$(2.3 \pm 1.0) \%$	
$\Gamma_{109}$	$\phi \rho^+$	$< 1.5$	$\%$	CL=90%
$\Gamma_{110}$	$K^+ K^- \pi^+ \pi^0$ non- $\phi$		$(1.5 \pm_{-0.6}^{+0.7}) \%$	
$\Gamma_{111}$	$K^*(892)^+ K_S^0$		$(1.7 \pm 0.8) \%$	

## Doubly Cabibbo-suppressed modes

$\Gamma_{112}$	$K^+ \pi^0$		$(1.89 \pm 0.25) \times 10^{-4}$	S=1.2
$\Gamma_{113}$	$K^+ \eta$		$(1.12 \pm 0.18) \times 10^{-4}$	
$\Gamma_{114}$	$K^+ \eta'(958)$		$(1.83 \pm 0.23) \times 10^{-4}$	
$\Gamma_{115}$	$K^+ \pi^+ \pi^-$		$(5.46 \pm 0.25) \times 10^{-4}$	
$\Gamma_{116}$	$K^+ \rho^0$		$(2.1 \pm 0.5) \times 10^{-4}$	
$\Gamma_{117}$	$K^*(892)^0 \pi^+, K^*(892)^0 \rightarrow$ $K^+ \pi^-$		$(2.6 \pm 0.4) \times 10^{-4}$	
$\Gamma_{118}$	$K^+ f_0(980), f_0(980) \rightarrow$ $\pi^+ \pi^-$		$(4.9 \pm 2.9) \times 10^{-5}$	
$\Gamma_{119}$	$K_2^*(1430)^0 \pi^+, K_2^*(1430)^0 \rightarrow$ $K^+ \pi^-$		$(4.4 \pm 3.0) \times 10^{-5}$	
$\Gamma_{120}$	$K^+ \pi^+ \pi^-$ nonresonant	not seen		
$\Gamma_{121}$	$2K^+ K^-$		$(9.0 \pm 2.1) \times 10^{-5}$	

 $\Delta C = 1$  weak neutral current (C1) modes, or

## Lepton Family number (LF) or Lepton number (L) violating modes

$\Gamma_{122}$	$\pi^+ e^+ e^-$	C1	$< 1.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{123}$	$\pi^+ \phi, \phi \rightarrow e^+ e^-$	[e]	$(1.7 \pm_{-0.9}^{+1.4}) \times 10^{-6}$		
$\Gamma_{124}$	$\pi^+ \mu^+ \mu^-$	C1	$< 7.3$	$\times 10^{-8}$	CL=90%
$\Gamma_{125}$	$\pi^+ \phi, \phi \rightarrow \mu^+ \mu^-$	[e]	$(1.8 \pm 0.8) \times 10^{-6}$		
$\Gamma_{126}$	$\rho^+ \mu^+ \mu^-$	C1	$< 5.6$	$\times 10^{-4}$	CL=90%
$\Gamma_{127}$	$K^+ e^+ e^-$	[f]	$< 1.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{128}$	$K^+ \mu^+ \mu^-$	[f]	$< 4.3$	$\times 10^{-6}$	CL=90%
$\Gamma_{129}$	$\pi^+ e^+ \mu^-$	LF	$< 2.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{130}$	$\pi^+ e^- \mu^+$	LF	$< 3.6$	$\times 10^{-6}$	CL=90%
$\Gamma_{131}$	$K^+ e^+ \mu^-$	LF	$< 1.2$	$\times 10^{-6}$	CL=90%
$\Gamma_{132}$	$K^+ e^- \mu^+$	LF	$< 2.8$	$\times 10^{-6}$	CL=90%
$\Gamma_{133}$	$\pi^- 2e^+$	L	$< 1.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{134}$	$\pi^- 2\mu^+$	L	$< 2.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{135}$	$\pi^- e^+ \mu^+$	L	$< 2.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{136}$	$\rho^- 2\mu^+$	L	$< 5.6$	$\times 10^{-4}$	CL=90%
$\Gamma_{137}$	$K^- 2e^+$	L	$< 9$	$\times 10^{-7}$	CL=90%
$\Gamma_{138}$	$K^- 2\mu^+$	L	$< 1.0$	$\times 10^{-5}$	CL=90%
$\Gamma_{139}$	$K^- e^+ \mu^+$	L	$< 1.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{140}$	$K^*(892)^- 2\mu^+$	L	$< 8.5$	$\times 10^{-4}$	CL=90%

$\Gamma_{141}$	Unaccounted decay modes		$(50.2 \pm 0.9) \%$	S=1.1
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[a] The branching fraction for this mode may differ from the sum of the submodes that contribute to it, due to interference effects. See the relevant papers.

[b] These subfractions of the  $K^- 2\pi^+$  mode are uncertain: see the Particle Listings.

[c] Submodes of the  $D^+ \rightarrow K^- 2\pi^+ \pi^0$  and  $K_S^0 2\pi^+ \pi^-$  modes were studied by ANJOS 92C and COFFMAN 92B, but with at most 142 events for the first mode and 229 for the second – not enough for precise results. With nothing new for 18 years, we refer to our 2008 edition, Physics Letters **B667** 1 (2008), for those results.

[d] The unseen decay modes of the resonances are included.

[e] This is *not* a test for the  $\Delta C=1$  weak neutral current, but leads to the  $\pi^+ \ell^+ \ell^-$  final state.

[f] This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.

## CONSTRAINED FIT INFORMATION

An overall fit to 23 branching ratios uses 31 measurements and one constraint to determine 15 parameters. The overall fit has a  $\chi^2 = 33.4$  for 17 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

# Meson Particle Listings

$D^\pm$

x29	0									
x34	0	0								
x35	22	0	0							
x38	5	0	0	1						
x40	18	0	0	4	26					
x50	0	0	0	0	0	0				
x60	0	0	0	0	0	0	0			
x61	0	0	0	0	0	0	0	0		
x62	5	0	0	1	8	29	0	0	0	
x99	5	0	0	1	7	27	0	0	0	77
x95	5	0	0	1	69	26	0	0	0	8
x96	11	0	0	2	16	63	0	0	0	18
x112	3	0	0	1	5	17	0	0	0	5
x141	-86	-2	-17	-36	-20	-48	-20	-18	-10	-21
	x16	x29	x34	x35	x38	x40	x50	x60	x61	x62
x95	7									
x96	17	16								
x112	5	5	11							
x141	-19	-19	-32	-9						
	x89	x95	x96	x112						

## D<sup>+</sup> BRANCHING RATIOS

Some now-obsolete measurements have been omitted from these Listings.

### c-quark decays

#### $\Gamma(c \rightarrow e^+ \text{ anything})/\Gamma(c \rightarrow \text{ anything})$

For the Summary Table, we only use the average of  $e^+$  and  $\mu^+$  measurements from  $Z^0 \rightarrow c\bar{c}$  decays; see the second data block below.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.103±0.009<sup>+0.009</sup><sub>-0.008</sub></b>	378	<sup>1</sup> ABBIENDI	99K OPAL	$Z^0 \rightarrow c\bar{c}$

<sup>1</sup>ABBIENDI 99K uses the excess of right-sign over wrong-sign leptons opposite reconstructed  $D^*(2010)^+ \rightarrow D^0\pi^+$  decays in  $Z^0 \rightarrow c\bar{c}$ .

#### $\Gamma(c \rightarrow \mu^+ \text{ anything})/\Gamma(c \rightarrow \text{ anything})$

For the Summary Table, we only use the average of  $e^+$  and  $\mu^+$  measurements from  $Z^0 \rightarrow c\bar{c}$  decays; see the next data block.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.082±0.005 OUR AVERAGE</b>				
0.073±0.008±0.002	73	KAYIS-TOPAK.05	CHRS	$\nu_\mu$ emulsion
0.095±0.007 <sup>+0.014</sup> <sub>-0.013</sub>	2829	ASTIER	00D NOMD	$\nu_\mu\text{Fe} \rightarrow \mu^-\mu^+X$
0.090±0.007 <sup>+0.007</sup> <sub>-0.006</sub>	476	<sup>1</sup> ABBIENDI	99K OPAL	$Z^0 \rightarrow c\bar{c}$
0.086±0.017 <sup>+0.008</sup> <sub>-0.007</sub>	69	<sup>2</sup> ALBRECHT	92F ARG	$e^+e^- \approx 10\text{ GeV}$
0.078±0.009±0.012		ONG	88 MRK2	$e^+e^- 29\text{ GeV}$
0.078±0.015±0.02		BARTEL	87 JADE	$e^+e^- 34.6\text{ GeV}$
0.082±0.012 <sup>+0.02</sup> <sub>-0.01</sub>		ALTHOFF	84G TASS	$e^+e^- 34.5\text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.093±0.009±0.009	88	KAYIS-TOPAK.02	CHRS	See KAYIS-TOPAKSU 05
0.089±0.018±0.025		BARTEL	85J JADE	See BARTEL 87

<sup>1</sup>ABBIENDI 99K uses the excess of right-sign over wrong-sign leptons opposite reconstructed  $D^*(2010)^+ \rightarrow D^0\pi^+$  decays in  $Z^0 \rightarrow c\bar{c}$ .

<sup>2</sup>ALBRECHT 92F uses the excess of right-sign over wrong-sign leptons in a sample of events tagged by fully reconstructed  $D^*(2010)^+ \rightarrow D^0\pi^+$  decays.

#### $\Gamma(c \rightarrow \ell^+ \text{ anything})/\Gamma(c \rightarrow \text{ anything})$

This is an average (not a sum) of  $e^+$  and  $\mu^+$  measurements.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.096 ±0.004 OUR AVERAGE</b>				
0.0958±0.0042±0.0028	1828	<sup>1</sup> ABREU	00o DLPH	$Z^0 \rightarrow c\bar{c}$
0.095 ±0.006 <sup>+0.007</sup> <sub>-0.006</sub>	854	<sup>2</sup> ABBIENDI	99K OPAL	$Z^0 \rightarrow c\bar{c}$

<sup>1</sup>ABREU 00o uses leptons opposite fully reconstructed  $D^*(2010)^+$ ,  $D^+$ , or  $D^0$  mesons.

<sup>2</sup>ABBIENDI 99K uses the excess of right-sign over wrong-sign leptons opposite reconstructed  $D^*(2010)^+ \rightarrow D^0\pi^+$  decays in  $Z^0 \rightarrow c\bar{c}$ .

#### $\Gamma(c \rightarrow D^*(2010)^+ \text{ anything})/\Gamma(c \rightarrow \text{ anything})$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.255±0.015±0.008</b>	2371	<sup>1</sup> ABREU	00o DLPH	$Z^0 \rightarrow c\bar{c}$

<sup>1</sup>ABREU 00o uses slow pions opposite fully reconstructed  $D^*(2010)^+$ ,  $D^+$ , or  $D^0$  mesons as a signal of  $D^*(2010)^-$  production.

### Inclusive modes

#### $\Gamma(e^+ \text{ semileptonic})/\Gamma_{\text{total}}$

The sum of our  $\bar{K}^0e^+\nu_e$ ,  $\bar{K}^*(892)^0e^+\nu_e$ ,  $\pi^0e^+\nu_e$ ,  $\eta e^+\nu_e$ ,  $\rho^0e^+\nu_e$ , and  $\omega e^+\nu_e$  branching fractions is  $15.3 \pm 0.4\%$ .

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>16.07±0.30 OUR AVERAGE</b>				
16.13±0.10±0.29	26.2±0.2k	<sup>1</sup> ASNER	10 CLEO	$e^+e^-$ at 3774 MeV
15.2 ±0.9 ±0.8	521 ± 32	ABLIKIM	07G BES2	$e^+e^- \approx \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
16.13±0.20±0.33	8798 ±105	<sup>2</sup> ADAM	06A CLEO	See ASNER 10
17.0 ±1.9 ±0.7	158	BALTRUSAIT...	85B MRK3	$e^+e^- 3.77\text{ GeV}$

<sup>1</sup>Using the  $D^+$  and  $D^0$  lifetimes, ASNER 10 finds that the ratio of the  $D^+$  and  $D^0$  semileptonic widths is  $0.985 \pm 0.015 \pm 0.024$ .

<sup>2</sup>Using the  $D^+$  and  $D^0$  lifetimes, ADAM 06A finds that the ratio of the  $D^+$  and  $D^0$  inclusive  $e^+$  widths is  $0.985 \pm 0.028 \pm 0.015$ , consistent with the isospin-invariance prediction of 1.

#### $\Gamma(\mu^+ \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>17.6±2.7±1.8</b>	100 ± 12	<sup>1</sup> ABLIKIM	08L BES2	$e^+e^- \approx \psi(3772)$

<sup>1</sup>ABLIKIM 08L finds the ratio of  $D^+ \rightarrow \mu^+X$  and  $D^0 \rightarrow \mu^+X$  branching fractions to be  $2.59 \pm 0.70 \pm 0.25$ , in accord with the ratio of  $D^+$  and  $D^0$  lifetimes,  $2.54 \pm 0.02$ .

#### $\Gamma(K^- \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>25.7±1.4 OUR AVERAGE</b>				
24.7±1.3±1.2	631 ± 33	ABLIKIM	07G BES2	$e^+e^- \approx \psi(3770)$
27.8 <sup>+3.6</sup> <sub>-3.1</sub>		BARLAG	92c ACCM	$\pi^- \text{ Cu } 230\text{ GeV}$
27.1±2.3±2.4		COFFMAN	91 MRK3	$e^+e^- 3.77\text{ GeV}$

#### $[\Gamma(\bar{K}^0 \text{ anything}) + \Gamma(K^0 \text{ anything})]/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>61 ±5 OUR AVERAGE</b>				
60.5±5.5±3.3	244 ± 22	ABLIKIM	06u BES2	$e^+e^-$ at 3773 MeV
61.2±6.5±4.3		COFFMAN	91 MRK3	$e^+e^- 3.77\text{ GeV}$

#### $\Gamma(K^+ \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>5.9±0.8 OUR AVERAGE</b>				
6.1±0.9±0.4	189 ± 27	ABLIKIM	07G BES2	$e^+e^- \approx \psi(3770)$
5.5±1.3±0.9		COFFMAN	91 MRK3	$e^+e^- 3.77\text{ GeV}$

#### $\Gamma(K^*(892)^- \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>5.7±5.2±0.7</b>	7.2 ± 6.5	ABLIKIM	06u BES2	$e^+e^-$ at 3773 MeV

#### $\Gamma(\bar{K}^*(892)^0 \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>23.2±4.5±3.0</b>	189 ± 36	ABLIKIM	05P BES	$e^+e^- \approx 3773\text{ MeV}$

#### $\Gamma(K^*(892)^0 \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.6</b>	90	ABLIKIM	05P BES	$e^+e^- \approx 3773\text{ MeV}$

#### $\Gamma(\eta \text{ anything})/\Gamma_{\text{total}}$

This ratio includes  $\eta$  particles from  $\eta'$  decays.

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>6.3±0.5±0.5</b>	1972 ±142	HUANG	06B CLEO	$e^+e^-$ at $\psi(3770)$

#### $\Gamma(\eta' \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1.04±0.16±0.09</b>	82 ± 13	HUANG	06B CLEO	$e^+e^-$ at $\psi(3770)$

#### $\Gamma(\phi \text{ anything})/\Gamma_{\text{total}}$

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1.03±0.10±0.07</b>	248 ± 21	HUANG	06B CLEO	$e^+e^-$ at $\psi(3770)$

### Leptonic and semileptonic modes

#### $\Gamma(e^+ \nu_e)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;8.8 × 10<sup>-6</sup></b>	90	EISENSTEIN	08 CLEO	$e^+e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.4 × 10 <sup>-5</sup>	90	ARTUSO	05A CLEO	See EISENSTEIN 08

$\Gamma(\mu^+ \nu_\mu)/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$   
See the note on "Decay Constants of Charged Pseudoscalar Mesons" in the  $D_S^+$  Listings.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.74 \pm 0.17</math> OUR AVERAGE</b>				
$3.71 \pm 0.19 \pm 0.06$	$409 \pm 21$	<sup>1</sup> ABLIKIM	14F BES3	$e^+e^-$ at $\psi(3770)$
$3.82 \pm 0.32 \pm 0.09$	$150 \pm 12$	<sup>2</sup> EISENSTEIN	08 CLEO	$e^+e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$12.2 \pm^{+11.1}_{-5.3} \pm 1.0$	3	<sup>3</sup> ABLIKIM	05D BES	$e^+e^- \approx 3.773$ GeV
$4.40 \pm 0.66 \pm^{+0.09}_{-0.12}$	$47 \pm 7$	<sup>4</sup> ARTUSO	05A CLEO	See EISENSTEIN 08
$3.5 \pm 1.4 \pm 0.6$	7	<sup>5</sup> BONVICINI	04A CLEO	Incl. in ARTUSO 05A
$8 \pm^{+16}_{-5} \pm^{+5}_{-2}$	1	<sup>6</sup> BAI	98B BES	$e^+e^- \rightarrow D^{*+} D^-$

<sup>1</sup> ABLIKIM 14F obtain  $|V_{cd}| \cdot f_{D^+} = (45.75 \pm 1.20 \pm 0.39)$  MeV, and using  $|V_{cd}| = 0.22520 \pm 0.00065$  gets  $f_{D^+} = (203.2 \pm 5.3 \pm 1.8)$  MeV.

<sup>2</sup> EISENSTEIN 08, using the  $D^+$  lifetime and assuming  $|V_{cd}| = |V_{us}|$ , gets  $f_{D^+} = (205.8 \pm 8.5 \pm 2.5)$  MeV from this measurement.

<sup>3</sup> ABLIKIM 05D finds a background-subtracted  $2.67 \pm 1.74$   $D^+ \rightarrow \mu^+ \nu_\mu$  events, and from this obtains  $f_{D^+} = 371 \pm^{+129}_{-119} \pm 25$  MeV.

<sup>4</sup> ARTUSO 05A obtains  $f_{D^+} = 222.6 \pm 16.7 \pm^{+2.8}_{-3.4}$  MeV from this measurement.

<sup>5</sup> BONVICINI 04A finds eight events with an estimated background of one, and from the branching fraction obtains  $f_{D^+} = 202 \pm 41 \pm 17$  MeV.

<sup>6</sup> BAI 98B obtains  $f_{D^+} = (300 \pm^{+180}_{-150} \pm^{+80}_{-40})$  MeV from this measurement.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.2 \times 10^{-3}</math></b>	90	EISENSTEIN	08 CLEO	$e^+e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.1 \times 10^{-3}$	90	RUBIN	06A CLEO	See EISENSTEIN 08

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>8.90 \pm 0.15</math> OUR AVERAGE</b>				
$8.962 \pm 0.054 \pm 0.206$	40k	<sup>1</sup> ABLIKIM	15AF BES3	from $D^+ \rightarrow K_L e^+ \nu_e$
$8.83 \pm 0.10 \pm 0.20$	8.5k	<sup>2</sup> BESSON	09 CLEO	from $D^+ \rightarrow K_S e^+ \nu_e$
$8.95 \pm 1.59 \pm 0.67$	34	<sup>3</sup> ABLIKIM	05A BES	from $D^+ \rightarrow K_S e^+ \nu_e$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.53 \pm 0.13 \pm 0.23$		<sup>4</sup> DOBBS	08 CLEO	See BESSON 09
$8.71 \pm 0.38 \pm 0.37$	545	HUANG	05B CLEO	See DOBBS 08

<sup>1</sup> ABLIKIM 15AF report  $\Gamma(D^+ \rightarrow K_L e^+ \nu_e)/\Gamma_{\text{total}} = (4.481 \pm 0.027 \pm 0.103)\%$ . See also the form-factor parameters near the end of this  $D^+$  Listing.

<sup>2</sup> See the form-factor parameters near the end of this  $D^+$  Listing.

<sup>3</sup> The ABLIKIM 05A result together with the  $D^0 \rightarrow K^- e^+ \nu_e$  branching fraction of ABLIKIM 04c and Particle Data Group lifetimes gives  $\Gamma(D^0 \rightarrow K^- e^+ \nu_e) / \Gamma(D^+ \rightarrow \bar{K}^0 e^+ \nu_e) = 1.08 \pm 0.22 \pm 0.07$ ; isospin invariance predicts the ratio is 1.0.

<sup>4</sup> DOBBS 08 establishes  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_{D^+}^{\pi}(0)}{f_{D^+}^{K}(0)}| = 0.188 \pm 0.008 \pm 0.002$  from the  $D^+$  and  $D^0$  decays to  $\bar{K} e^+ \nu_e$  and  $\pi e^+ \nu_e$ . It also finds  $\Gamma(D^0 \rightarrow K^- e^+ \nu_e) / \Gamma(D^+ \rightarrow \bar{K}^0 e^+ \nu_e) = 1.06 \pm 0.02 \pm 0.03$ ; isospin invariance predicts the ratio is 1.0.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.093 \pm 0.007</math> OUR FIT</b>				
<b><math>0.103 \pm 0.023 \pm 0.008</math></b>	$29 \pm 6$	ABLIKIM	07 BES2	$e^+e^-$ at 3773 MeV

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.99 \pm 0.07</math> OUR FIT</b>				
<b><math>1.019 \pm 0.076 \pm 0.065</math></b>	$555 \pm 39$	LINK	04E FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.50 \pm 0.75 \pm 0.27$	$29 \pm 6$	ABLIKIM	06o BES2	$e^+e^-$ at 3773 MeV
$3.5 \pm^{+1.2}_{-0.7} \pm 0.4$	14	BAI	91 MRK3	$e^+e^- \approx 3.77$ GeV

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.4380 \pm 0.0036 \pm 0.0042</math></b>	$70k \pm 363$	DEL-AMO-SA...11i	BABR	$e^+e^- \approx 10.6$ GeV

$\Gamma(\bar{K}^*(892)^0 e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{34}/\Gamma$   
Unseen decay modes of  $\bar{K}^*(892)^0$  are included. See the end of the  $D^+$  Listings for measurements of  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$  form-factor ratios.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.52 \pm 0.15</math> OUR FIT</b>				
<b><math>5.52 \pm 0.07 \pm 0.13</math></b>	$\approx 5k$	BRIERE	10 CLEO	$e^+e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$5.06 \pm 1.21 \pm 0.40$	$28 \pm 7$	ABLIKIM	06o BES2	$e^+e^-$ at 3773 MeV
$5.56 \pm 0.27 \pm 0.23$	$422 \pm 21$	<sup>1</sup> HUANG	05B CLEO	$e^+e^-$ at $\psi(3770)$
<sup>1</sup> HUANG 05B finds $\Gamma(D^0 \rightarrow K^{*-} e^+ \nu_e) / \Gamma(D^+ \rightarrow \bar{K}^{*0} e^+ \nu_e) = 0.98 \pm 0.08 \pm 0.04$ ; isospin invariance predicts the ratio is 1.0.				

$\Gamma(\bar{K}^*(892)^0 e^+ \nu_e)/\Gamma(K^- 2\pi^+)$   $\Gamma_{34}/\Gamma_{40}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included. See the end of the  $D^+$  Listings for measurements of  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$  form-factor ratios.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.74 \pm 0.04 \pm 0.05$		BRANDENB...	02 CLEO	$e^+e^- \approx \gamma(4S)$
$0.62 \pm 0.15 \pm 0.09$	35	ADAMOVICH	91 OMEG	$\pi^-$ 340 GeV
$0.55 \pm 0.08 \pm 0.10$	880	ALBRECHT	91 ARG	$e^+e^- \approx 10.4$ GeV
$0.49 \pm 0.04 \pm 0.05$		ANJOS	89B E691	Photoproduction

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>94.11 \pm 0.74 \pm 0.75</math></b>	DEL-AMO-SA...11i	BABR	$e^+e^- \approx 10.6$ GeV

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>5.79 \pm 0.16 \pm 0.15</math></b>	DEL-AMO-SA...11i	BABR	$e^+e^- \approx 10.6$ GeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;6 \times 10^{-3}</math></b>	90	DEL-AMO-SA...11i	BABR	$e^+e^- \approx 10.6$ GeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5 \times 10^{-4}</math></b>	90	DEL-AMO-SA...11i	BABR	$e^+e^- \approx 10.6$ GeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.007</math></b>	90	ANJOS	89B E691	Photoproduction

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.417 \pm 0.030 \pm 0.023</math></b>	$555 \pm 39$	LINK	04E FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.30 \pm 0.15</math> OUR FIT</b>				
<b><math>5.27 \pm 0.07 \pm 0.14</math></b>	$\approx 5k$	BRIERE	10 CLEO	$e^+e^-$ at $\psi(3770)$

$\Gamma(\bar{K}^*(892)^0 \mu^+ \nu_\mu)/\Gamma(K^- 2\pi^+)$   $\Gamma_{35}/\Gamma_{16}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included. See the end of the  $D^+$  Listings for measurements of  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$  form-factor ratios.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.57 \pm 0.04</math> OUR FIT</b>				Error includes scale factor of 1.1.
<b><math>0.594 \pm 0.043 \pm 0.033</math></b>	$555 \pm 39$	LINK	04E FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\bar{K}^*(892)^0 \mu^+ \nu_\mu)/\Gamma(K^- 2\pi^+)$   $\Gamma_{35}/\Gamma_{40}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included. See the end of the  $D^+$  Listings for measurements of  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$  form-factor ratios.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.561 \pm 0.022</math> OUR FIT</b>				Error includes scale factor of 1.3.
<b><math>0.57 \pm 0.06</math> OUR AVERAGE</b>				Error includes scale factor of 1.2.
$0.72 \pm 0.10 \pm 0.05$		BRANDENB...	02 CLEO	$e^+e^- \approx \gamma(4S)$
$0.56 \pm 0.04 \pm 0.06$	875	FRABETTI	93E E687	$\gamma$ Be $\bar{E}_\gamma \approx 200$ GeV
$0.46 \pm 0.07 \pm 0.08$	224	KODA MA	92C E653	$\pi^-$ emulsion 600 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.602 \pm 0.010 \pm 0.021$	12k	<sup>1</sup> LINK	02J FOCS	$\gamma$ nucleus, $\approx 180$ GeV

<sup>1</sup> This LINK 02J result includes the effects of an interference of a small S-wave  $K^- \pi^+$  amplitude with the dominant  $\bar{K}^{*0}$  amplitude. (The interference effect is reported in LINK 02E.) This result is redundant with results of LINK 04E elsewhere in these Listings.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.0530 \pm 0.0074 \pm^{+0.0099}_{-0.0096}</math></b>	14k	LINK	05i FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.042</math></b>	90	FRABETTI	93E E687	$\gamma$ Be $\bar{E}_\gamma \approx 200$ GeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.0064</math></b>	90	LINK	05i FOCS	$\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.04</math></b>	90	LINK	05i FOCS	$\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

# Meson Particle Listings

$D^\pm$

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.405±0.016±0.009</b>	838	<sup>1</sup> BESSON 09	CLEO	$e^+ e^-$ at $\psi(3770)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.373±0.022±0.013		<sup>2</sup> DOBBS 08	CLEO	See BESSON 09	
0.44 ±0.06 ±0.03	63 ± 9	HUANG 05B	CLEO	See DOBBS 08	

<sup>1</sup> See the form-factor parameters near the end of this  $D^+$  Listing.

<sup>2</sup> DOBBS 08 establishes  $|\frac{V_{cs}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}| = 0.188 \pm 0.008 \pm 0.002$  from the  $D^+$  and  $D^0$  decays to  $\overline{K} e^+ \nu_e$  and  $\pi e^+ \nu_e$ . It finds  $\Gamma(D^0 \rightarrow \pi^- e^+ \nu_e) / \Gamma(D^+ \rightarrow \pi^0 e^+ \nu_e) = 2.03 \pm 0.14 \pm 0.08$ ; isospin invariance predicts the ratio is 2.0.

$\Gamma(\eta e^+ \nu_e)/\Gamma_{\text{total}}$				$\Gamma_{28}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>11.4±0.9±0.4</b>		YELTON 11	CLEO	$e^+ e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13.3±2.0±0.6	46 ± 8	MITCHELL 09B	CLEO	See YELTON 11

$\Gamma(\rho^0 e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>2.18±<sup>+0.17</sup><sub>-0.25</sub> OUR FIT</b>						

<b>2.17 ± 0.12<sup>+0.12</sup><sub>-0.22</sub></b>	447 ± 25	<sup>1</sup> DOBBS	13	CLEO	$e^+e^-$ at $\psi(3770)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
2.1 ± 0.4 ± 0.1	27 ± 6	<sup>2</sup> HUANG	05B	CLEO	See DOBBS 13	
<sup>1</sup> DOBBS 13 finds $\Gamma(D^0 \rightarrow \rho^- e^+ \nu_e) / 2 \Gamma(D^+ \rightarrow \rho^0 e^+ \nu_e) = 1.03 \pm 0.09^{+0.08}_{-0.02}$ ; isospin invariance predicts the ratio is 1.0.						
<sup>2</sup> HUANG 05B finds $\Gamma(D^0 \rightarrow \rho^- e^+ \nu_e) / 2 \Gamma(D^+ \rightarrow \rho^0 e^+ \nu_e) = 1.2^{+0.4}_{-0.3} \pm 0.1$ ; isospin invariance predicts the ratio is 1.0.						

$\Gamma(\rho^0 e^+ \nu_e)/\Gamma(\bar{K}^*(892)^0 e^+ \nu_e)$					$\Gamma_{29}/\Gamma_{34}$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>0.0396^{+0.0033}_{-0.0050}</math> OUR FIT</b>						
<b><math>0.045 \pm 0.014 \pm 0.009</math></b>	49	<sup>1</sup> AITALA	97	E791	$\pi^-$ nucleus, 500 GeV	
<sup>1</sup> AITALA 97 explicitly subtracts $D^+ \rightarrow \eta' e^+ \nu_e$ and other backgrounds to get this result.						

$(\rho^0 \mu^+ \nu_\mu)/\Gamma(\bar{K}^*(892)^0 \mu^+ \nu_\mu)$					$\Gamma_{30}/\Gamma_{35}$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>0.045±0.007 OUR AVERAGE</b>		Error includes scale factor of 1.1.				
0.041±0.006±0.004	320 ± 44	LINK	06B	FOCS	$\gamma$ A, $\overline{E}_\gamma \approx 180$ GeV	
0.051±0.015±0.009	54	<sup>1</sup> AITALA	97	E791	$\pi^-$ nucleus, 500 GeV	
0.079±0.019±0.013	39	<sup>2</sup> FRABETTI	97	E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV	

<sup>1</sup> AITALA 97 explicitly subtracts  $D^+ \rightarrow \eta' \mu^+ \nu_\mu$  and other backgrounds to get this result.  
<sup>2</sup> Because the reconstruction efficiency for photons is low, this FRABETTI 97 result also includes any  $D^+ \rightarrow \eta' \mu^+ \nu_\mu \rightarrow \gamma \rho^0 \mu^+ \nu_\mu$  events in the numerator.

$\Gamma(\omega e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.69±0.11 OUR AVERAGE</b>					
1.63±0.11±0.08	491 ± 32	ABLIKIM 15W	BES3	292 fb <sup>-1</sup> , 3773 MeV	
1.82±0.18±0.07	129 ± 13	DOBBS 13	CLEO	$e^+ e^-$ at $\psi(3770)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.6 $\begin{smallmatrix} +0.7 \\ -0.6 \end{smallmatrix}$ ±0.1	7.6 $\begin{smallmatrix} +3.3 \\ -2.7 \end{smallmatrix}$	HUANG 05B	CLEO	See DOBBS 13	

$\Gamma(\eta'(958) e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{32}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.16 \pm 0.53 \pm 0.07</math></b>		YELTON 11	CLEO	$e^+ e^-$ at $\psi(3770)$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
<3.5	90	MITCHELL 09B	CLEO	See YELTON 11	

$\Gamma(\phi e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$	
Unseen decay modes of the $\phi$ are included.						
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;1.3 <math>\times 10^{-5}</math></b>	90	ABLIKIM	15W	BES3	292 fb <sup>-1</sup> , 3773 MeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.9 $\times 10^{-4}$	90	YELTON	11	CLEO	$e^+ e^-$ at $\psi(3770)$	
<1.6 $\times 10^{-4}$	90	MITCHELL	09B	CLEO	See YELTON 11	
<0.0201	90	ABLIKIM	06P	BES2	$e^+ e^-$ at 3773 MeV	
<0.0209	90	BAI	91	MRK3	$e^+ e^- \approx 3.77$ GeV	

## Hadronic modes with a $\overline{K}$ or $\overline{K} K \overline{K}$

$\Gamma(K_S^0 \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{38}/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>1.53 <math>\pm</math> 0.06 OUR FIT</b>	Error includes scale factor of 2.8.					
<b>1.578 <math>\pm</math> 0.013 <math>\pm</math> 0.025</b>		BONVICINI	14	CLEO	All CLEO-c runs	

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.526±0.022±0.038		<sup>1</sup> DOBBS	07	CLEO	See MENDEZ 10
1.55 ±0.05 ±0.06	2.2k	<sup>1</sup> HE	05	CLEO	See DOBBS 07
1.6 ±0.3 ±0.1	161	ADLER	88c	MRK3	e <sup>+</sup> e <sup>-</sup> 3.77 GeV
<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.					

$\Gamma(K_S^0 \pi^+)/\Gamma(K^- 2\pi^+)$					$\Gamma_{38}/\Gamma_{40}$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>0.162 ±0.007 OUR FIT</b>		Error includes scale factor of 3.3.				
<b>0.1530 ±0.0023 ±0.0016</b>	10.6k	LINK	02b	FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.1682 ±0.0012 ±0.0037	30k	MENDEZ	10	CLEO	See BONVICINI 14	
0.174 ±0.012 ±0.011	473	<sup>1</sup> BISHAI	97	CLEO	$e^+ e^- \approx 7(45)$	
0.137 ±0.015 ±0.016	264	ANJOS	90c	E691	Photoproduction	

$\Gamma(K_L^0 \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{39}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.460±0.040±0.035</b>	2023 ± 54	<sup>1</sup> HE 08	CLEO	$e^+e^-$ at $\psi(3770)$	
<sup>1</sup> The difference of CLEO $D^+ \rightarrow K_S^0 \pi^+$ and $K_L^0 \pi^+$ branching fractions over the sum (DOBBS 07 and HE 08) is $+0.022 \pm 0.016 \pm 0.018$ .					

$\Gamma(K^- 2\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{40}/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>9.46 ±0.24 OUR FIT</b>	Error includes scale factor of 2.0.					
<b>9.224±0.059±0.157</b>		BONVICINI	14	CLEO	All CLEO-c runs	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
9.14 ±0.10 ±0.17		<sup>1</sup> DOBBS	07	CLEO	See BONVICINI 14	
9.5 ±0.2 ±0.3	15.1k	<sup>1</sup> HE	05	CLEO	See DOBBS 07	
9.3 ±0.6 ±0.8	1502	<sup>2</sup> BALEST	94	CLEO	$e^+e^- \approx 7(45)$	
6.4 <sup>+1.5</sup> <sub>-1.4</sub>		<sup>3</sup> BARLAG	92c	ACCM	$\pi^-$ Cu 230 GeV	
9.1 ±1.3 ±0.4	1164	ADLER	88c	MRK3	$e^+e^-$ 3.77 GeV	
9.1 ±1.9	239	<sup>4</sup> SCHINDLER	81	MRK2	$e^+e^-$ 3.771 GeV	

<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.  
<sup>2</sup> BALEST 94 measures the ratio of  $D^+ \rightarrow K^- \pi^+ \pi^+$  and  $D^0 \rightarrow K^- \pi^+$  branching fractions to be  $2.35 \pm 0.16 \pm 0.16$  and uses their absolute measurement of the  $D^0 \rightarrow K^- \pi^+$  fraction (AKERIB 93).  
<sup>3</sup> BARLAG 92c computes the branching fraction by topological normalization.  
<sup>4</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.38 \pm 0.05$  nb. We use the MARK-3 (ADLER 88c) value of  $\sigma = 4.2 \pm 0.6 \pm 0.3$  nb.

## REVIEW OF MULTIBODY CHARM ANALYSES

Revised 2015 by D. M. Asner (Pacific Northwest National Laboratory) and J. Rademacker (University of Bristol)

**Kinematics & Models** The differential decay rate to a point  $s = (s_1, \dots, s_n)$  in  $n$  dimensional phase space can be expressed as

$$d\Gamma = |\mathcal{M}(s)|^2 \left| \frac{\partial^n \phi}{\partial(s_1 \dots s_n)} \right| d^n s \tag{1}$$

where  $|\frac{\partial^n \phi}{\partial(s_1 \dots s_n)}|$  represents the density of states at  $s$ , and  $\mathcal{M}$  the matrix element for the decay at that point in phase space. For two-body decays,  $|\frac{\partial^n \phi}{\partial(s_1 \dots s_n)}|$  is a  $\delta$  function, while for  $D^0$  decays to 3,4,5, . . . pseudoscalars, phase space is 2,5,8, . . . dimensional, leading to a rich phenomenology. Additional parameters are required to fully describe decays with vector particles in the initial or final state.

For the important case of a pseudoscalar decaying to 3 pseudoscalars, the decay kinematics can be described in a two dimensional Dalitz plot [1]. The Dalitz plot of  $D \rightarrow abc$  is usually parametrized in terms of invariant-mass-squared variables  $s_1 = (p_a + p_b)^2$  and  $s_2 = (p_b + p_c)^2$ , where  $p_a, p_b, p_c$  are the four-momenta of particles  $a, b, c$ . In terms of these variables, phase-space density is constant across the kinematically allowed region, so that any structure seen in the Dalitz plot is a direct consequence of the dynamics encoded in  $|\mathcal{M}|^2$ .



An important difference between decays to two or three pseudoscalars compared to decays to four or more particles is the behavior under parity. In the former case, the operation of parity can also be expressed as a rotation, so no parity violating observables can be defined (unless they also violate rotational invariance). This is not the case for decays to four or more particles. This leads to the interesting possibility of using parity-odd observables in four body decays for  $CP$  violation searches, as discussed below. Another consequence of these considerations is that four-body-decay kinematics cannot be described unambiguously in terms of invariant-mass-squared variables, as these are all parity even.

The matrix element  $\mathcal{M}$  is usually modeled as a sum of interfering decay amplitudes, each proceeding through resonant two-body decays [2]. See Refs [2–4] for a review of resonance phenomenology. In most analyses, each resonance is described by a Breit–Wigner or Flatté lineshape, and the model includes a non-resonant term with a constant phase and magnitude across the Dalitz plot. This approach has well-known theoretical limitations, such as the violation of unitarity and analyticity, which tend to be particularly problematic for broad, overlapping resonances. This motivates the use of more sophisticated descriptions, especially for the broad, overlapping resonances that occur typically in the S-wave components. In charm analyses, these have included the K-matrix approach [5,6,7] which respects unitarity; the use of LASS scattering data [8]; dispersive methods [9,10]; methods based on chiral symmetry [11,12]; and quasi model-independent parametrizations [13,14]. An important example first analyzed by CLEO [15,16,17] is  $D^0 \rightarrow K_S \pi^+ \pi^-$ , which is a key channel in  $CP$  violation and charm mixing analyses. Belle models this final state as a superposition of 18 resonances (including 4 significant doubly Cabibbo suppressed amplitudes) described by Breit–Wigner or Flatté lineshapes, plus a non-resonant component [18]. CDF’s analysis follows a similar approach [19]. BaBar’s model for the same decay replaces the broad  $\pi\pi$  and  $K\pi$  S-wave resonances and the non-resonant component with a K-matrix description [20]. Belle’s and BaBar’s data have been re-analyzed by [21] in a QCD factorization framework, using line-shape parametrizations for the S [11,12] and P wave [10] contributions (with input from  $\tau^- \rightarrow K_S \pi^- \nu_\tau$  data [22] for the latter) that preserve 2-body unitarity and analyticity. The measurements give compatible results for the components they share. All three approaches remain within the confines of the “isobar” framework which treats the decay as a series of independent two-body processes, ignoring long-range hadronic effects. Dispersive techniques that account for these hadronic effects and respect full 3 body unitarity and analyticity have been applied to regions of the  $D^- \rightarrow K^- \pi^+ \pi^+$  Dalitz plot below the  $\eta'K$  threshold [23].

Limitations in the theoretical description of interfering resonances are the leading source of systematic uncertainty in many analyses. This is set to become increasingly problematic given the statistical precision achievable with the vast charm samples

available at the B factories, LHCb, and their upgrades. Already now, clean data samples with millions of charm events are available even in suppressed decay modes, e.g. 2.4M  $D^0 \rightarrow \pi^- \pi^+ \pi^0$  events at LHCb [24]. In some cases, the model uncertainty can be removed through model-independent amplitude methods, often relying on input from the charm threshold, as discussed below. At the same time, increasingly sophisticated models are being developed, and applied to data.

**Applications of multibody charm analyses** The interference between the decay paths via which multibody decays proceed provides sensitivity to both relative magnitudes and phases of the contributing decay amplitudes. It is especially this sensitivity to phases that makes amplitude analyses such a uniquely powerful tool for studying a wide range of phenomena. Here we concentrate on their use for  $CP$  violation measurements and mixing in charm, and charm inputs to  $CP$  violation analyses in  $B$  meson decays. The properties of light-meson resonances determined in  $D$ -meson amplitude analyses are reported in the light-unflavored-meson section of this *Review*.

#### **Time-integrated searches for $CP$ violation in charm**

Comparing the results of amplitude fits for  $CP$ -conjugate decay modes provides a measure of  $CP$  violation. The advantage of this approach over the model-independent searches discussed in the next paragraph is the physical interpretation of any  $CP$  violation observation that such a fit result would allow. The disadvantage lies in the theoretical uncertainty intrinsic to such analyses due to the amplitude-model dependence. Recent  $CP$  violation searches using this method include CLEO’s amplitude analysis of  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$  [25] and CDF’s analysis of  $\sim 350,000$   $D^0 \rightarrow K_S \pi^+ \pi^-$  events [19].

The most common model-independent approach for searching for local  $CP$  violation across a Dalitz plot is based on performing a  $\chi^2$  comparison of the number of events in the bins of  $CP$ -conjugate Dalitz plots. This method was pioneered by BaBar [26] and developed further in [27,28], with recent results in  $D^\pm \rightarrow K^+ K^- \pi^\pm$  [29,30,31],  $D^0 \rightarrow K_S \pi^+ \pi^-$  [19], and  $D^+ \rightarrow \pi^- \pi^+ \pi^+$  [32]. These techniques have been generalized to four-body decays, and applied to  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$  and  $D^0 \rightarrow \pi^+ \pi^- \pi^+ \pi^-$  [33]. Un-binned methods can increase the sensitivity [34]; two different unbinned methods have been applied by LHCb to  $D^+ \rightarrow \pi^- \pi^+ \pi^+$  [32] and  $D^0 \rightarrow \pi^+ \pi^- \pi^0$  [24]. None of these analyses have shown evidence of  $CP$  violation.

Another model-independent approach, providing complementary information, is based on constructing observables in four body decays that are odd under motion reversal (“naïve T”) [35–43], which is equivalent to  $P$  for scalar particles [43]. One such observable is  $C_T = \vec{p}_{K^+} \cdot (\vec{p}_{\pi^+} \times \vec{p}_{\pi^-})$  in  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$ . The rate asymmetry of positive and negative  $C_T$ ,  $A_T \equiv \frac{\Gamma(C_T > 0) - \Gamma(C_T < 0)}{\Gamma(C_T > 0) + \Gamma(C_T < 0)}$ , is a  $P$  violating parameter. Comparing  $A_T$  with the  $C$ -conjugate asymmetry in  $\overline{D}^0$  decays,  $\overline{A}_T$ , provides sensitivity to  $CP$  violation. Searches for  $CP$  violation in this manner have been carried out by FOCUS in  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$  [44], BaBar in  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$ ,

# Meson Particle Listings

$D^\pm$

$D^+ \rightarrow K^+ K_S \pi^+ \pi^-$ , and  $D_s^+ \rightarrow K^+ K_S \pi^+ \pi^-$  [45,46], and LHCb in  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$  [47]. In addition to a phase-space integrated result, LHCb's analysis is also carried out locally in sub-regions of phase space to enhance the sensitivity of the method. All results so far have been consistent with  $CP$  conservation.

**$D$  mixing and  $CP$  violation** Time-dependent amplitude analyses in decays to final states that are accessible to both  $D^0$  and  $\bar{D}^0$  have unique sensitivity to mixing parameters. A Dalitz plot analysis of a self-conjugate final state, such as  $K_S \pi^+ \pi^-$  and  $K_S K^+ K^-$ , allows the measurement of the phase difference between the relevant  $D^0$  and  $\bar{D}^0$  decay amplitudes, and thus a direct measurement of the mixing parameters  $x, y$  (rather than the decay-specific parameters  $x'^2, y'$  measured for example in  $D^0 \rightarrow K\pi$ ) [17]. These analyses are also sensitive to  $CP$  violation in mixing and in the interference between mixing and decay. These results are summarized in Ref. [48]. The important role from charm threshold data as input to such measurements is discussed below.

**Charm amplitude analyses for measuring  $\gamma/\phi_3$**  Neutral  $D$  mesons originating from  $B^- \rightarrow DK^-$  (which we denote with  $D_{B^-}$ ) are a superposition of  $D^0$  and  $\bar{D}^0$  with a relative phase that depends on  $\gamma/\phi_3$ :

$$D_{B^-} \propto D^0 + r_B e^{i(\delta_B - \gamma)} \bar{D}^0,$$

where  $\delta_B$  is a  $CP$  conserving strong phase, and  $r_B \sim 0.1$ . In the corresponding  $CP$ -conjugate expression,  $\gamma/\phi_3$  changes sign. An amplitude analysis of the subsequent decay of the  $D_{B^\pm}$  allows an extraction of  $\gamma/\phi_3$  [49–54]. The method generalizes to similar  $B$  hadron decays, such as  $B^0 \rightarrow DK^{*0}$ . Measurements based on this technique have been reported by BaBar, Belle and LHCb using both model-dependent approaches and model-independent ones based on CLEO-c input [18,55–61,65–67]. The most precise individual results come from the study of the  $D_{B^-} \rightarrow K_S \pi^+ \pi^-$  and  $D_{B^-} \rightarrow K_S K^+ K^-$  with an uncertainty of approximately  $15^\circ$  [18,55,59,67].

## Model independent methods for $\gamma/\phi_3$ and charm mixing

The theoretical uncertainty on amplitude models of multibody  $D^0$  decays potentially limits the precision of measurements of  $\gamma/\phi_3$  in  $B^\pm \rightarrow DK^\pm$  and related decay modes. Model-independent methods to measure  $\gamma/\phi_3$  require input related to the relative phases of the  $D^0$  and  $\bar{D}^0$  decay amplitudes across the phase-space distribution. The same considerations apply to measurements of  $D^0$  mixing and  $CP$  violation parameters in time-dependent Dalitz plot analyses. The required phase information is accessible at the charm threshold, where CLEO-c and BES III operate [48,52,68–74]. There,  $D$  mesons originate from the decay  $\psi(3770) \rightarrow D\bar{D}$ . The two  $D$  mesons are quantum-correlated which can be used to identify decays of well-defined  $D^0 - \bar{D}^0$  superpositions to the final state of interest. The resulting interference of  $D^0$  and  $\bar{D}^0$  amplitudes provides the desired model-independent phase information. For decays to non-self-conjugate decays such as  $D^0 \rightarrow K^+ \pi^- \pi^+ \pi^-$ , analysing  $D^0 - \bar{D}^0$

superpositions provides the only way of measuring the relative phase between the  $D^0$  and  $\bar{D}^0$  amplitudes.

These analyses can be performed in sub-regions/bins of phase space, or integrated across phase space. The relevant result can be expressed in terms of one complex parameter  $\mathcal{Z} = R e^{-i\delta}$  per pair of  $CP$ -conjugate phase space bins, with magnitude  $R \leq 1$ . The larger  $R$ , the higher the sensitivity to interference effects, and thus to  $\gamma/\phi_3$ . The sensitivity of the binned analyses can be optimized by using amplitude model-dependent information to maximize  $R$  in each bin, without introducing a model-dependent bias in the result. CLEO-c data have been analyzed in this way to provide binned results for the self-conjugate decays  $D^0 \rightarrow K_S \pi\pi$  and  $D^0 \rightarrow K_S K K$  [75,76]. The phase-space integrated analyses for  $D^0, \bar{D}^0 \rightarrow K_S K^+ \pi^-$ ,  $K^- \pi^+ \pi^0$ , and  $K^- \pi^+ \pi^- \pi^+$  have yielded  $\mathcal{Z}_{K_S K \pi} = (0.73 \pm 0.8) e^{-i(8.3^\circ \pm 15.2^\circ)}$ ,  $\mathcal{Z}_{K \pi \pi^0} = (0.82 \pm 0.07) e^{-i(164^\circ \pm 20^\circ)}$ ,  $\mathcal{Z}_{K 3\pi} = (0.32_{-0.28}^{+0.20}) e^{-i(225^\circ \pm 21^\circ)}$ , respectively [77,78,79]. These results follow the usual convention for  $\gamma/\phi_3$ -related studies where  $CP|D^0\rangle = +|\bar{D}^0\rangle$ , while in charm mixing measurements, one usually takes  $CP|D^0\rangle = -|\bar{D}^0\rangle$ , leading to a phase-shift in  $\delta$  of  $\pi$ . Restricting the analysis to a bin around the  $K^* K$  resonance in the  $K_S K \pi$  Dalitz plot, [77] find  $R = 1.00 \pm 0.16$ , illustrating the benefit in dividing phase space into bins.

The corresponding phase space-integrated input for self-conjugate decays such as  $D^0 \rightarrow \pi^+ \pi^- \pi^0$  takes the form of a single real parameter, the  $CP$ -even fraction  $F_+$ , defined such that a  $CP$  even eigenstate has  $F_+ = 1$ , while a  $CP$ -odd eigenstate has  $F_+ = 0$  [72]. A recent analysis of CLEO-c data revealed that  $D^0 \rightarrow \pi^+ \pi^- \pi^0$  is compatible with being completely  $CP$ -even with  $F_+ = 1.014 \pm 0.045 \pm 0.022$ , while  $D^0 \rightarrow K^+ K^- \pi^0$  has  $F_+ = 0.734 \pm 0.106 \pm 0.054$  and  $D^0 \rightarrow \pi^+ \pi^- \pi^+ \pi^-$  has  $F_+ = 0.737 \pm 0.028$  [73].

The charm system itself provides, through mixing, a well-defined, time-dependent superposition of  $D^0$  and  $\bar{D}^0$ . Using mixing parameters measured independently as input, this can be used to obtain the relevant information for  $\gamma/\phi_3$  measurements. This method is expected to be particularly powerful in doubly Cabibbo-suppressed decays such as  $D^0 \rightarrow K^+ \pi^- \pi^+ \pi^-$ , and when used in conjunction with information from charm threshold [80,81].

**Summary** Multibody charm decays offer a rich phenomenology, including unique sensitivity to  $CP$  violation and charm mixing. This is a highly dynamic field with many new results (some of which we presented here) and rapidly increasing, high quality datasets. These datasets constitute a huge opportunity, but also a challenge to improve the theoretical descriptions of soft hadronic effects in multibody decays. For some measurements, model-independent methods, many relying on input from the charm threshold, provide a way of removing model-induced uncertainties. At the same time, work is ongoing to improve the theoretical description of multibody decays.

## References

1. R.H. Dalitz, *Phil. Mag.* **44**, 1068 (1953).

2. M. Bauer, B. Stech, and M. Wirbel, Z. Phys. **C4**, 103 (1987); P. Bedaque, A. Das, and V.S. Mathur, Phys. Rev. **D49**, 269 (1994); L.-L. Chau and H.-Y. Cheng, Phys. Rev. **D36**, 137 (1987); K. Terasaki, Int. J. Mod. Phys. **A10**, 3207 (1995); F. Buccella, M. Lusignoli, and A. Pugliese, Phys. Lett. **B379**, 249 (1996).
3. J.D. Jackson, Nuovo Cimento **34**, 1644 (1964).
4. See the note on Kinematics in this *Review*.
5. E.P. Wigner, Phys. Rev. **70**, 15 (1946).
6. S.U. Chung *et al.*, Ann. Phys. **4**, 404 (1995).
7. I.J.R. Aitchison, Nucl. Phys. **A189**, 417 (1972).
8. D. Aston *et al.* (LASS Collab.), Nucl. Phys. **B296**, 493 (1988).
9. R. Omnes, Nuovo Cimento **8**, 316 (1958).
10. C. Hanhart, Phys. Lett. **B715**, 170 (2012).
11. B. El-Bennich *et al.* Phys. Rev. **D79**, 094005 (2009) [Phys. Rev. **D83**, 039903(E) (2011)].
12. J.P. Dedonder *et al.* Acta Phys. Polon. B **42** (2011) 2013.
13. E.M. Aitala *et al.* (E791 Collab.), Phys. Rev. **D73**, 032004 (2006) [Phys. Rev. **D74**, 059901 (E)(2006)].
14. G. Bonvicini *et al.* (CLEO Collab.), Phys. Rev. **D78**, 052001 (2008).
15. H. Muramatsu *et al.* (CLEO Collab.), Phys. Rev. Lett. **89**, 251802 (2002).
16. D.M. Asner *et al.* (CLEO Collab.), Phys. Rev. **D70**, 091101 (2004).
17. D.M. Asner *et al.* (CLEO Collab.), Phys. Rev. **D72**, 012001 (2005).
18. A. Poluektov *et al.* (Belle Collab.), Phys. Rev. **D81**, 112002 (2010).
19. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D86**, 032007 (2012).
20. P. del Amo Sanchez *et al.* (BaBar Collab.), Phys. Rev. Lett. **105**, 081803 (2010).
21. J.P. Dedonder *et al.* Phys. Rev. **D89**, 094018 (2014).
22. D. Epifanov *et al.* (Belle Collab.), Phys. Lett. **B654**, 65 (2007).
23. F. Niecknig and B. Kubis, JHEP **1510**, 142 (2015).
24. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B740**, 158 (2015).
25. M. Artuso *et al.* (CLEO Collab.), Phys. Rev. **D85**, 122002 (2012).
26. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D78**, 051102 (2008).
27. I. Bediaga *et al.*, Phys. Rev. **D80**, 096006 (2009).
28. I. Bediaga *et al.*, Phys. Rev. **D86**, 036005 (2012).
29. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D87**, 052010 (2013).
30. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. **D84**, 112008 (2011).
31. R. Aaij *et al.* (LHCb Collab.), JHEP **1306**, 112 (2013).
32. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B728**, 585 (2014).
33. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B726**, 623 (2013).
34. M. Williams, Phys. Rev. **D84**, 054015 (2011).
35. E. Golowich and G. Valencia, Phys. Rev. **D40**, 112, (1989).
36. G. Valencia, Phys. Rev. **D39**, 3339 (1989).
37. W. Bensalem and D. London, Phys. Rev. **D64**, 116003 (2001).
38. I.I.Y. Bigi, hep-ph/0107102.
39. W. Bensalem, A. Datta, and D. London, Phys. Rev. **D66**, 094004 (2002).
40. W. Bensalem, A. Datta, and D. London, Phys. Lett. **B538**, 309 (2002).
41. A. Datta and D. London, Int. J. Mod. Phys. **A19**, 2505 (2004).
42. M. Gronau and J.L. Rosner, Phys. Rev. **D84**, 096013 (2011).
43. G. Durieux and Y. Grossman, Phys. Rev. **D92**, 076013 (2015).
44. J.M. Link *et al.* (FOCUS Collab.), Phys. Lett. **B622**, 239 (2005).
45. P. del Amo Sanchez *et al.* (BaBar Collab.), Phys. Rev. **D81**, 111103 (2010).
46. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D84**, 031103 (2011).
47. R. Aaij *et al.* (LHCb Collab.), JHEP **1410**, 005 (2014).
48. See the note on  $D^0-\bar{D}^0$  Mixing in this *Review*.
49. M. Gronau and D. Wyler, Phys. Lett. **B265**, 172 (1991).
50. M. Gronau and D. London, Phys. Lett. **B253**, 483 (1991).
51. D. Atwood *et al.*, Phys. Rev. Lett. **78**, 17, 3257 (1997).
52. A. Giri *et al.*, Phys. Rev. **D68**, 5, 054018 (2003).
53. A. Poluektov *et al.* (Belle Collab.), Phys. Rev. **D70**, 7, 072003 (2004).
54. J. Rademacker and G. Wilkinson, Phys. Lett. **B647**, 400 (2007).
55. P. del Amo Sanchez *et al.* (BaBar Collab.), Phys. Rev. Lett. **105**, 121801 (2010).
56. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B726**, 151 (2013).
57. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B723**, 44 (2013).
58. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B718**, 43 (2012).
59. R. Aaij *et al.* (LHCb Collab.), JHEP **1410**, 97 (2014).
60. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B733**, 36 (2014).
61. R. Aaij *et al.* (LHCb Collab.), Nucl. Phys. **B888**, 169 (2014).
62. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **99**, 251801 (2007).
63. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D84**, 012002 (2011).
64. M. Nayak *et al.* (Belle Collab.), Phys. Rev. **D88**, 091104 (2013).
65. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. **D91**, 112014 (2015).
66. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D84**, 012002 (2011).
67. H. Aihara *et al.* (Belle Collab.), Phys. Rev. **D85**, 112014 (2012).
68. D. Atwood and A. Soni, Phys. Rev. **D68**, 033003 (2003).
69. S. Malde and G. Wilkinson, Phys. Lett. **B701**, 353 (2011).
70. A. Bondar *et al.*, Phys. Rev. **D82**, 034033 (2010).
71. C. Thomas and G. Wilkinson, JHEP **1210**, 184 (2012).

## Meson Particle Listings

 $D^\pm$ 

72. M. Nayak *et al.* Phys. Lett. **B740**, 1 (2015).  
 73. S. Malde *et al.* Phys. Lett. **B747**, 9 (2015).  
 74. S. Malde, C. Thomas, and G. Wilkinson, Phys. Rev. **D91**, 094032 (2015).  
 75. J. Libby *et al.* (CLEO Collab.), Phys. Rev. **D82**, 112006 (2010).  
 76. R.A. Briere *et al.* (CLEO Collab.), Phys. Rev. **D80**, 032002 (2009).  
 77. J. Insler *et al.* (CLEO Collab.), Phys. Rev. **D85**, 092016 (2012).  
 78. J. Libby *et al.*, Phys. Lett. **B731**, 197 (2014).  
 79. N. Lowrey *et al.* (CLEO Collab.), Phys. Rev. **D80**, 031105 (2009).  
 80. S. Harnew and J. Rademacker, Phys. Lett. **B728**, 296 (2014).  
 81. S. Harnew and J. Rademacker, JHEP **1503**, 169 (2015).

 $\Gamma((K^-\pi^+)_{S\text{-wave}}\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{41}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis. The  $K^-\pi^+$  S-wave includes a broad scalar  $\kappa$  ( $\bar{K}_0^*(800)$ ), the  $\bar{K}_0^*(1430)^0$ , and non-resonant background.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.801 ± 0.012 OUR AVERAGE</b>			
0.8024 ± 0.0138 ± 0.0043	<sup>1</sup> LINK	09	FOCS MIPWA fit, 53k evts
0.838 ± 0.038	<sup>2</sup> BONVICINI	08A	CLEO QMIPWA fit, 141k evts
0.786 ± 0.014 ± 0.018	AITALA	06	E791 Dalitz fit, 15.1k events
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.8323 ± 0.0150 ± 0.0008	<sup>3</sup> LINK	07B	FOCS See LINK 09

<sup>1</sup> This LINK 09 model-independent partial-wave analysis of the  $K^-\pi^+$  S-wave slices the  $K^-\pi^+$  mass range into 39 bins.

<sup>2</sup> The BONVICINI 08A QMIPWA (quasi-model-independent partial-wave analysis) of the  $K^-\pi^+$  S-wave amplitude slices the  $K^-\pi^+$  mass range into 26 bins but keeps the Breit-Wigner  $\bar{K}_0^*(1430)^0$ .

<sup>3</sup> This LINK 07B fit uses a K matrix. The  $K^-\pi^+$  S-wave fit fraction given above breaks down into  $(207.3 \pm 25.5 \pm 12.4)\%$  isospin-1/2 and  $(40.5 \pm 9.6 \pm 3.2)\%$  isospin-3/2 — with large interference between the two. The isospin-1/2 component includes the  $\kappa$  (or  $\bar{K}_0^*(800)^0$ ) and  $\bar{K}_0^*(1430)^0$ .

 $\Gamma(\bar{K}_0^*(800)^0\pi^+, \bar{K}_0^*(800) \rightarrow K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{42}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.478 ± 0.121 ± 0.053	AITALA	02	E791 See AITALA 06

 $\Gamma(\bar{K}^*(892)^0\pi^+, \bar{K}^*(892)^0 \rightarrow K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{44}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.111 ± 0.012 OUR AVERAGE</b>			Error includes scale factor of 3.7.
0.1236 ± 0.0034 ± 0.0034	LINK	09	FOCS MIPWA fit, 53k evts
0.0988 ± 0.0046	BONVICINI	08A	CLEO QMIPWA fit, 141k evts
0.119 ± 0.002 ± 0.020	AITALA	06	E791 Dalitz fit, 15.1k events
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1361 ± 0.0041 ± 0.0030	<sup>1</sup> LINK	07B	FOCS See LINK 09
0.123 ± 0.010 ± 0.009	AITALA	02	E791 See AITALA 06
0.137 ± 0.006 ± 0.009	FRABETTI	94G	E687 Dalitz fit, 8800 evts
0.170 ± 0.009 ± 0.034	ANJOS	93	E691 $\gamma$ Be 90–260 GeV
0.14 ± 0.04 ± 0.04	ALVAREZ	91B	NA14 Photoproduction
0.13 ± 0.01 ± 0.07	ADLER	87	MRK3 $e^+e^-$ 3.77 GeV

<sup>1</sup> The statistical error on this LINK 07B value is corrected in LINK 09.

 $\Gamma(\bar{K}^*(1410)^0\pi^+, \bar{K}^{*0} \rightarrow K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{45}/\Gamma_{40}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>not seen</b>	LINK	09	FOCS MIPWA fit, 53k evts
<b>not seen</b>	BONVICINI	08A	CLEO QMIPWA fit, 141k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.8 ± 2.1 ± 1.7	LINK	07B	FOCS See LINK 09

 $\Gamma(\bar{K}_0^*(1430)^0\pi^+, \bar{K}_0^*(1430)^0 \rightarrow K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{43}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

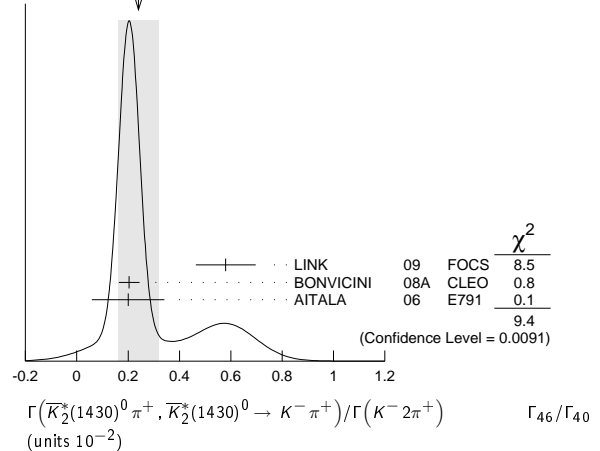
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.1330 ± 0.0062</b>	BONVICINI	08A	CLEO QMIPWA fit, 141k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.125 ± 0.014 ± 0.005	AITALA	02	E791 See AITALA 06
0.284 ± 0.022 ± 0.059	FRABETTI	94G	E687 Dalitz fit, 8800 evts
0.248 ± 0.019 ± 0.017	ANJOS	93	E691 $\gamma$ Be 90–260 GeV

 $\Gamma(\bar{K}_2^*(1430)^0\pi^+, \bar{K}_2^*(1430)^0 \rightarrow K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{46}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.24 ± 0.08 OUR AVERAGE</b>			Error includes scale factor of 2.2. See the ideogram below.
0.58 ± 0.10 ± 0.06	LINK	09	FOCS MIPWA fit, 53k evts
0.204 ± 0.040	BONVICINI	08A	CLEO QMIPWA fit, 141k evts
0.2 ± 0.1 ± 0.1	AITALA	06	E791 Dalitz fit, 15.1k events
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.39 ± 0.09 ± 0.05	LINK	07B	FOCS See LINK 09
0.5 ± 0.1 ± 0.2	AITALA	02	E791 See AITALA 06

WEIGHTED AVERAGE  
0.24 ± 0.08 (Error scaled by 2.2)


 $\Gamma(\bar{K}^*(1680)^0\pi^+, \bar{K}^*(1680)^0 \rightarrow K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{47}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.23 ± 0.12 OUR AVERAGE</b>			
1.75 ± 0.62 ± 0.54	LINK	09	FOCS MIPWA fit, 53k evts
0.196 ± 0.118	BONVICINI	08A	CLEO QMIPWA fit, 141k evts
1.2 ± 0.6 ± 1.2	AITALA	06	E791 Dalitz fit, 15.1k events
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.90 ± 0.63 ± 0.43	LINK	07B	FOCS See LINK 09
2.5 ± 0.7 ± 0.3	AITALA	02	E791 See AITALA 06
4.7 ± 0.6 ± 0.7	FRABETTI	94G	E687 Dalitz fit, 8800 evts
3.0 ± 0.4 ± 1.3	ANJOS	93	E691 $\gamma$ Be 90–260 GeV

 $\Gamma(K^-(2\pi^+)_{I=2})/\Gamma(K^-2\pi^+)$   $\Gamma_{48}/\Gamma_{40}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.155 ± 0.028</b>	BONVICINI	08A	CLEO QMIPWA fit, 141k evts

 $\Gamma(K^-2\pi^+ \text{ nonresonant})/\Gamma(K^-2\pi^+)$   $\Gamma_{49}/\Gamma_{40}$ 

This is the “fit fraction” from the Dalitz-plot analysis. Later analyses find little need for this decay mode.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.130 ± 0.058 ± 0.044	AITALA	02	E791 See AITALA 06
0.998 ± 0.037 ± 0.072	FRABETTI	94G	E687 Dalitz fit, 8800 evts
0.838 ± 0.088 ± 0.275	ANJOS	93	E691 $\gamma$ Be 90–260 GeV
0.79 ± 0.07 ± 0.15	ADLER	87	MRK3 $e^+e^-$ 3.77 GeV

 $\Gamma(K_S^0\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{50}/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.24 ± 0.17 OUR FIT</b>				
<b>7.244 ± 0.053 ± 0.166</b>		BONVICINI	14	CLEO All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.99 ± 0.09 ± 0.25	<sup>1</sup> DOBBS	07	CLEO	See BONVICINI 14
7.2 ± 0.2 ± 0.4	<sup>1</sup> HE	05	CLEO	See DOBBS 07
5.1 ± 1.3 ± 0.8	159	ADLER	88c	MRK3 $e^+e^-$ 3.77 GeV

<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.

 $\Gamma(K_S^0\rho^+)/\Gamma(K_S^0\pi^+\pi^0)$   $\Gamma_{51}/\Gamma_{50}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>83.4 ± 2.2 ± 7.1 / 3.6</b>	<sup>1</sup> ABLIKIM	14E	BES3 $e^+e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
68 ± 8 ± 12	ADLER	87	MRK3 $e^+e^-$ 3.77 GeV

<sup>1</sup> Fit fraction from Dalitz plot analysis of 142k  $D^+ \rightarrow K_S^0\pi^+\pi^0$  events.

$\Gamma(K_S^0 \rho(1450)^+, \rho^+ \rightarrow \pi^+ \pi^0) / \Gamma(K_S^0 \pi^+ \pi^0)$	$\Gamma_{52} / \Gamma_{50}$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$2.1 \pm 0.3^{+1.6}_{-1.9}$	ABLIKIM	14E	BES3 $e^+ e^-$ at $\psi(3770)$

$\Gamma(\bar{K}^*(892)^0 \pi^+, \bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0) / \Gamma(K_S^0 \pi^+ \pi^0)$			$\Gamma_{53} / \Gamma_{50}$
This is the “fit fraction” from the Dalitz-plot analysis.			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
$3.58 \pm 0.17^{+0.39}_{-0.38}$	<sup>1</sup> ABLIKIM	14E	BES3 $e^+ e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
19 $\pm 6$ $\pm 6$	ADLER	87	MRK3 $e^+ e^-$ 3.77 GeV
<sup>1</sup> Fit fraction from Dalitz plot analysis of 142k $D^+ \rightarrow K_S^0 \pi^+ \pi^0$ events.			

$\Gamma(\bar{K}_0^*(1430)^0 \pi^+, \bar{K}_0^* \rightarrow K_S^0 \pi^0) / \Gamma(K_S^0 \pi^+ \pi^0)$				$\Gamma_{54} / \Gamma_{50}$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
$3.7 \pm 0.6 \pm 1.1$	ABLIKIM	14E	BES3	$e^+ e^-$ at $\psi(3770)$

$\Gamma(\bar{K}_0^*(1680)^0 \pi^+, \bar{K}_0^* \rightarrow K_S^0 \pi^0) / \Gamma(K_S^0 \pi^+ \pi^0)$	$\Gamma_{55} / \Gamma_{50}$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$1.3 \pm 0.2^{+0.9}_{-1.3}$	ABLIKIM	14E	BES3 $e^+ e^-$ at $\psi(3770)$

$\Gamma(\pi^0 \pi^+, \pi^0 \rightarrow K_S^0 \pi^0)/\Gamma(K_S^0 \pi^+ \pi^0)$	$\Gamma_{56}/\Gamma_{50}$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$7.7 \pm 1.2^{+6.5}_{-4.8}$	ABLIKIM	14E	BES3 $e^+ e^-$ at $\psi(3770)$

$\Gamma(K_S^0 \pi^+ \pi^0 \text{ nonresonant})/\Gamma(K_S^0 \pi^+ \pi^0)$	$\Gamma_{57}/\Gamma_{50}$		
This is the “fit fraction” from the Dalitz-plot analysis.			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
$4.6 \pm 0.7^{+5.4}_{-5.1}$	<sup>1</sup> ABLIKIM	14E	BES3 $e^+ e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$13 \pm 7 \pm 8$	ADLER	87	MRK3 $e^+ e^-$ 3.77 GeV
<sup>1</sup> Fit fraction from Dalitz plot analysis of 142k $D^+ \rightarrow K_S^0 \pi^+ \pi^0$ events.			

$\Gamma(K_S^0 \pi^+ \pi^0 \text{ nonresonant and } \pi^0 \pi^+) / \Gamma(K_S^0 \pi^+ \pi^0)$	$\Gamma_{58} / \Gamma_{50}$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$18.6 \pm 1.7^{+2.3}_{-4.6}$	ABLIKIM	14E BES3	$e^+ e^-$ at $\psi(3770)$

$\Gamma((K_S^0 \pi^0)_{S\text{-wave}} \pi^+) / \Gamma(K_S^0 \pi^+ \pi^0)$	$\Gamma_{59} / \Gamma_{50}$		
The numerator here is the coherent sum of the $\bar{K}_0^*(1430)^0 \pi^+$ , $\pi^0 \pi^+$ , and nonresonant contributions.			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$17.3 \pm 1.4^{+3.4}_{-4.3}$	ABLIKIM	14E BES3	$e^+ e^-$ at $\psi(3770)$

$\Gamma(K^-2\pi^+\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{60}/\Gamma$			
See our 2008 Review (Physics Letters <b>B667</b> 1 (2008)) for measurements of submodes of this mode. There is nothing new since 1992, and the two papers, ANJOS 92c, with $91 \pm 12$ events above background, and COFFMAN 92b, with $142 \pm 20$ such events, could not determine submode fractions with much accuracy.				
$\frac{VALUE (units\ 10^{-2})}{\underline{\hspace{1cm}}}$	$\frac{EVTS}{\underline{\hspace{1cm}}}$	$\frac{DOCUMENT\ ID}{\underline{\hspace{1cm}}}$	$\frac{TECN}{\underline{\hspace{1cm}}}$	$\frac{COMMENT}{\underline{\hspace{1cm}}}$
<b>6.14 <math>\pm 0.16</math> OUR FIT</b>				
<b>6.142 <math>\pm 0.045 \pm 0.154</math></b>		BONVICINI	14	CLEO All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.98 $\pm 0.08 \pm 0.16$		<sup>1</sup> DOBBS	07	CLEO See BONVICINI 14
6.0 $\pm 0.2 \pm 0.2$	4.8k	<sup>1</sup> HE	05	CLEO See DOBBS 07
5.8 $\pm 1.2 \pm 1.2$	142	COFFMAN	92b	MRK3 $e^+e^-$ 3.77 GeV
6.3 $\pm 1.4^{+1.4}_{-1.3} \pm 1.2$	175	BALTRUSAIT..86E	MRK3	See COFFMAN 92b

<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.

$\Gamma(K_S^0 2\pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{61}/\Gamma$			
See our 2008 Review (Physics Letters <b>B667</b> 1 (2008)) for measurements of submodes of this mode. There is nothing new since 1992, and the two papers, ANJOS 92c, with $229 \pm 17$ events above background, and COFFMAN 92b, with $209 \pm 20$ such events, could not determine submode fractions with much accuracy.				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.05 <math>\pm 0.09</math> OUR FIT</b>				
<b>3.051 <math>\pm 0.027 \pm 0.082</math></b>		BONVICINI	14	CLEO All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.122 $\pm 0.046 \pm 0.096$		<sup>1</sup> DOBBS	07	CLEO See BONVICINI 14
3.2 $\pm 0.1$ $\pm 0.2$	3.2k	<sup>1</sup> HE	05	CLEO See DOBBS 07
2.1 $\pm 1.0$ $-0.9$		<sup>2</sup> BARLAG	92c	ACCM $\pi^-$ Cu 230 GeV
3.3 $\pm 0.8$ $\pm 0.2$	168	ADLER	88c	MRK3 $e^+ e^-$ 3.77 GeV

<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.

<sup>2</sup> BARLAG 92c computes the branching fraction by topological normalization.

$\Gamma(K^- 3\pi^+ \pi^-)/\Gamma(K^- 2\pi^+)$					$\Gamma_{62}/\Gamma_{40}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.061<math>\pm</math>0.005 OUR FIT</b>		Error includes scale factor of 1.1.			
<b>0.062<math>\pm</math>0.008 OUR AVERAGE</b>		Error includes scale factor of 1.3.			
0.058 $\pm$ 0.002 $\pm$ 0.006	2923	LINK	03b	FOCS	$\gamma$ A, $\overline{E}_\gamma \approx 180$ GeV
0.077 $\pm$ 0.008 $\pm$ 0.010	239	FRABETTI	97c	E687	$\gamma$ Be, $\overline{E}_\gamma \approx 200$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.09 $\pm$ 0.01 $\pm$ 0.01	113	ANJOS	90b	E691	Photoproduction

$\Gamma(\bar{K}^*(892)^0 2\pi^+ \pi^-, \bar{K}^*(892)^0 \rightarrow K^- \pi^+) / \Gamma(K^- 3\pi^+ \pi^-)$	$\Gamma_{63} / \Gamma_{62}$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.21 ± 0.04 ± 0.06</b>	LINK	03D	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \pi^+, \bar{K}^*(892)^0 \rightarrow K^- \pi^+) / \Gamma(K^- 3\pi^+ \pi^-)$				$\Gamma_{64} / \Gamma_{62}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.40 \pm 0.03 \pm 0.06</math></b>	LINK	03b	FOCS	$\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \pi^+, \bar{K}^*(892)^0 \rightarrow K^- \pi^+) / \Gamma(K^- 2\pi^+)$				$\Gamma_{64} / \Gamma_{40}$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.016 \pm 0.007 \pm 0.004$	FRABETTI	97c	E687 $\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV	

$\Gamma(\bar{K}^*(892)^0 2\pi^+ \pi^- \text{ no-}\rho, \bar{K}^*(892)^0 \rightarrow K^- \pi^+) / \Gamma(K^- 2\pi^+)$	$\Gamma_{66} / \Gamma_{40}$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.032 \pm 0.010 \pm 0.008$	FRABETTI	97c	E687 $\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$\Gamma(K-\rho^0 2\pi^+)/\Gamma(K-3\pi^+ \pi^-)$	$\Gamma_{67}/\Gamma_{62}$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.30±0.04±0.01</b>	LINK	03b	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(K-\rho^0 2\pi^+)/\Gamma(K-2\pi^+)$	$\Gamma_{67}/\Gamma_{40}$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.034 \pm 0.009 \pm 0.005$	FRABETTI	97c	E687 $\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$\Gamma(\bar{K}^*(892)^0 a_1(1260)^+) / \Gamma(K^- 2\pi^+)$				$\Gamma_{65} / \Gamma_{40}$
Unseen decay modes of the $\bar{K}^*(892)^0$ and $a_1(1260)^+$ are included.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.099 ± 0.008 ± 0.018</b>	LINK	03D	FOCS	$\gamma$ A, $E_\gamma \approx 180$ GeV

$\Gamma(K^- 3\pi^+ \pi^- \text{ nonresonant})/\Gamma(K^- 3\pi^+ \pi^-)$				$\Gamma_{68}/\Gamma_{62}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.07 <math>\pm 0.05 \pm 0.01</math></b>		LINK	03D	FOCS $\gamma$ A, $\overline{E}_\gamma \approx 180$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.026	90	FRABETTI	97c	E687 $\gamma$ Be, $\overline{E}_\gamma \approx 200$ GeV

$\Gamma(K^+ 2K_S^0) / \Gamma(K^- 2\pi^+)$				$\Gamma_{69} / \Gamma_{40}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.049 ± 0.022 OUR AVERAGE</b>		Error includes scale factor of 2.4.		
0.035 ± 0.010 ± 0.005	39 ± 9	ALBRECHT	94i	ARG $e^+ e^- \approx 10$ GeV
0.085 ± 0.018	70 ± 12	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV

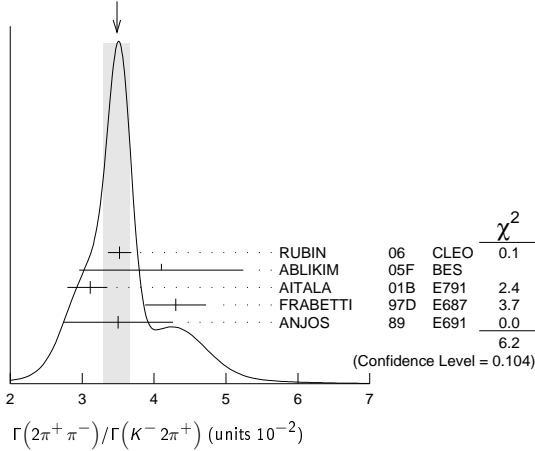
$\Gamma(K^+ K^- K_S^0 \pi^+)/\Gamma(K_S^0 2\pi^+ \pi^-)$					$\Gamma_{70}/\Gamma_{61}$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>7.7 \pm 1.5 \pm 0.9</math></b>	$35 \pm 7$	LINK	01c	FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV

#### Pionic modes

$\Gamma(\pi^+\pi^0)/\Gamma(K^-2\pi^+)$					$\Gamma_{71}/\Gamma_{40}$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.31<math>\pm</math>0.06 OUR AVERAGE</b>					
1.29 $\pm$ 0.04 $\pm$ 0.05	2649 $\pm$ 76	MENDEZ	10	CLEO $e^+e^-$ at 3774 MeV	
1.33 $\pm$ 0.11 $\pm$ 0.09	1229 $\pm$ 99	AUBERT,B	06f	BABR $e^+e^- \approx 7(45)$	
1.44 $\pm$ 0.19 $\pm$ 0.10	171 $\pm$ 22	ARMS	04	CLEO $e^+e^- \approx 10$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.33 $\pm$ 0.07 $\pm$ 0.06	914 $\pm$ 46	RUBIN	06	CLEO See MENDEZ 10	

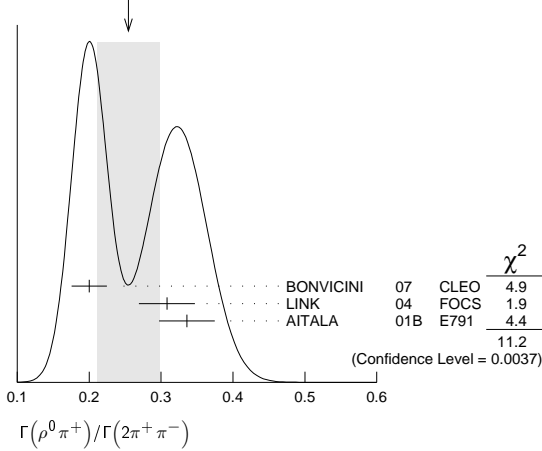
$\Gamma(2\pi^+\pi^-)/\Gamma(K^-2\pi^+)$				$\Gamma_{72}/\Gamma_{40}$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.48<math>\pm</math>0.19 OUR AVERAGE</b> Error includes scale factor of 1.4. See the ideogram below.				
3.52 $\pm$ 0.11 $\pm$ 0.12	3303 $\pm$ 95	RUBIN	06 CLEO	$e^+e^-$ at $\psi(3770)$
4.1 $\pm$ 1.1 $\pm$ 0.3	85 $\pm$ 22	ABLIKIM	05F BES	$e^+e^- \approx \psi(3770)$
3.11 $\pm$ 0.18 $^{+0.16}_{-0.26}$	1172	AITALA	01B E791	$\pi^-$ nucleus, 500 GeV
4.3 $\pm$ 0.3 $\pm$ 0.3	236	FRABETTI	97D E687	$\gamma$ Be $\approx$ 200 GeV
3.5 $\pm$ 0.7 $\pm$ 0.3	83	ANJOS	89 E691	Photoproduction

## Meson Particle Listings

 $D^\pm$ WEIGHTED AVERAGE  
3.48±0.19 (Error scaled by 1.4) $\Gamma(\rho^0 \pi^+)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{73}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.25 ± 0.04 OUR AVERAGE</b>			Error includes scale factor of 2.4. See the ideogram below.
0.200 ± 0.023 ± 0.009	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
0.3082 ± 0.0314 ± 0.0230	LINK 04	FOCS	Dalitz fit, 1527 ± 51 evts
0.336 ± 0.032 ± 0.022	AITALA 01B	E791	Dalitz fit, 1172 evts

WEIGHTED AVERAGE  
0.25±0.04 (Error scaled by 2.4) $\Gamma(\pi^+ (\pi^+ \pi^-)_{S\text{-wave}})/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{74}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis. See also the next three data blocks.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.5600 ± 0.0324 ± 0.0214</b>	<sup>1</sup> LINK 04	FOCS	Dalitz fit, 1527 ± 51 evts

<sup>1</sup> LINK 04 borrows a K-matrix parametrization from ANISOVICH 03 of the full  $\pi\pi$ -S-wave isoscalar scattering amplitude to describe the  $\pi^+\pi^-$  S-wave component of the  $\pi^+\pi^+\pi^-$  state. The fit fraction given above is a sum over five  $f_0$  mesons, the  $f_0(980)$ ,  $f_0(1300)$ ,  $f_0(1200\text{--}1600)$ ,  $f_0(1500)$ , and  $f_0(1750)$ . See LINK 04 for details and discussion.

 $\Gamma(\sigma \pi^+, \sigma \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{75}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.422 ± 0.027 OUR AVERAGE</b>			
0.418 ± 0.014 ± 0.025	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
0.463 ± 0.090 ± 0.021	AITALA 01B	E791	Dalitz fit, 1172 evts

 $\Gamma(f_0(980) \pi^+, f_0(980) \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{76}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.048 ± 0.010 OUR AVERAGE</b>			Error includes scale factor of 1.3.
0.041 ± 0.009 ± 0.003	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
0.062 ± 0.013 ± 0.004	AITALA 01B	E791	Dalitz fit, 1172 evts

 $\Gamma(f_0(1370) \pi^+, f_0(1370) \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{77}/\Gamma_{72}$ 

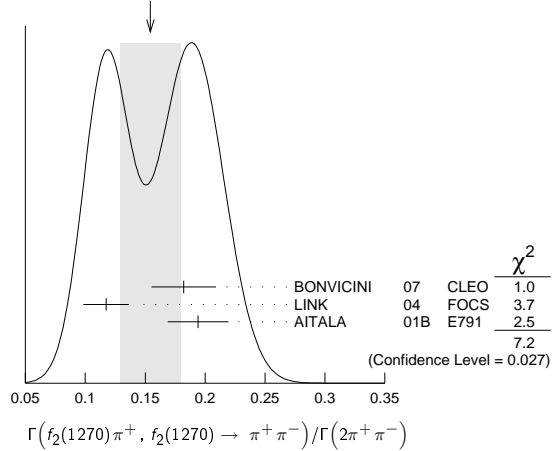
This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.024 ± 0.013 OUR AVERAGE</b>			
0.026 ± 0.018 ± 0.006	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
0.023 ± 0.015 ± 0.008	AITALA 01B	E791	Dalitz fit, 1172 evts

 $\Gamma(f_2(1270) \pi^+, f_2(1270) \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{78}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.154 ± 0.025 OUR AVERAGE</b>			Error includes scale factor of 1.9. See the ideogram below.
0.182 ± 0.026 ± 0.007	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
0.1174 ± 0.0190 ± 0.0029	LINK 04	FOCS	Dalitz fit, 1527 ± 51 evts
0.194 ± 0.025 ± 0.004	AITALA 01B	E791	Dalitz fit, 1172 evts

WEIGHTED AVERAGE  
0.154±0.025 (Error scaled by 1.9) $\Gamma(\rho(1450)^0 \pi^+, \rho(1450)^0 \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{79}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.024</b>	95	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.007 ± 0.007 ± 0.003		AITALA 01B	E791	Dalitz fit, 1172 evts

 $\Gamma(f_0(1500) \pi^+, f_0(1500) \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{80}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.034 ± 0.010 ± 0.008</b>	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts

 $\Gamma(f_0(1710) \pi^+, f_0(1710) \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{81}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.016</b>	95	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts

 $\Gamma(f_0(1790) \pi^+, f_0(1790) \rightarrow \pi^+ \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{82}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.02</b>	95	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts

 $\Gamma((\pi^+ \pi^+)_{S\text{-wave}} \pi^-)/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{83}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.037</b>	95	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts

 $\Gamma(2\pi^+ \pi^- \text{ nonresonant})/\Gamma(2\pi^+ \pi^-)$  $\Gamma_{84}/\Gamma_{72}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.035</b>	95	BONVICINI 07	CLEO	Dalitz fit, $\approx 2240$ evts
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.078 ± 0.060 ± 0.027		AITALA 01B	E791	Dalitz fit, 1172 evts

 $\Gamma(\pi^+ 2\pi^0)/\Gamma(K^- 2\pi^+)$  $\Gamma_{85}/\Gamma_{40}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.0 ± 0.3 ± 0.3</b>	1535 ± 89	RUBIN 06	CLEO	$e^+ e^-$ at $\psi(3770)$

 $\Gamma(2\pi^+ \pi^- \pi^0)/\Gamma(K^- 2\pi^+)$  $\Gamma_{86}/\Gamma_{40}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>12.4 ± 0.5 ± 0.6</b>	5701 ± 205	RUBIN 06	CLEO	$e^+ e^-$ at $\psi(3770)$

 $\Gamma(\eta \pi^+)/\Gamma_{\text{total}}$  $\Gamma_{90}/\Gamma$ Unseen decay modes of the  $\eta$  are included.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>34.3 ± 1.4 ± 1.7</b>	1033 ± 42	ARTUSO 08	CLEO	See MENDEZ 10

See key on page 601

## Meson Particle Listings

 $D^\pm$  $\Gamma(\eta\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{90}/\Gamma_{40}$ Unseen decay modes of the  $\eta$  are included.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.87 \pm 0.09 \pm 0.19</math></b>	2940 $\pm$ 68	MENDEZ	10	CLEO $e^+e^-$ at 3774 MeV
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$3.81 \pm 0.26 \pm 0.21$	377 $\pm$ 26	RUBIN	06	CLEO See ARTUSO 08

 $\Gamma(\omega\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{92}/\Gamma$ Unseen decay modes of the  $\omega$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.4 \times 10^{-4}</math></b>	90	RUBIN	06	CLEO $e^+e^-$ at $\psi(3770)$

 $\Gamma(3\pi^+2\pi^-)/\Gamma(K^-2\pi^+)$   $\Gamma_{89}/\Gamma_{40}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.77 \pm 0.17</math> OUR FIT</b>				
<b><math>1.73 \pm 0.20 \pm 0.17</math></b>	732 $\pm$ 77	RUBIN	06	CLEO $e^+e^-$ at $\psi(3770)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$2.3 \pm 0.4 \pm 0.2$	58	FRABETTI	97c	E687 $\gamma\text{Be}, \bar{E}_\gamma \approx 200$ GeV

 $\Gamma(3\pi^+2\pi^-)/\Gamma(K^-3\pi^+\pi^-)$   $\Gamma_{89}/\Gamma_{62}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.289 \pm 0.019</math> OUR FIT</b>				
<b><math>0.290 \pm 0.017 \pm 0.011</math></b>	835	LINK	03d	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

 $\Gamma(\eta\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{91}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>13.8 \pm 3.1 \pm 1.6</math></b>	149 $\pm$ 34	ARTUSO	08	CLEO $e^+e^-$ at $\psi(3770)$

 $\Gamma(\eta'(958)\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{93}/\Gamma$ Unseen decay modes of the  $\eta'(958)$  are included.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$44.2 \pm 2.5 \pm 2.9$	352 $\pm$ 20	ARTUSO	08	CLEO See MENDEZ 10

 $\Gamma(\eta'(958)\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{93}/\Gamma_{40}$ Unseen decay modes of the  $\eta'(958)$  are included.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.12 \pm 0.17 \pm 0.25</math></b>	1037 $\pm$ 35	MENDEZ	10	CLEO $e^+e^-$ at 3774 MeV

 $\Gamma(\eta'(958)\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{94}/\Gamma$ Unseen decay modes of the  $\eta'(958)$  are included.

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>15.7 \pm 4.3 \pm 2.5</math></b>	33 $\pm$ 9	ARTUSO	08	CLEO $e^+e^-$ at $\psi(3770)$

Hadronic modes with a  $K\bar{K}$  pair $\Gamma(K^+K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{95}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$3.14 \pm 0.09 \pm 0.08$	1971 $\pm$ 51	BONVICINI	08	CLEO See MENDEZ 10

 $\Gamma(K^+K_S^0)/\Gamma(K_S^0\pi^+)$   $\Gamma_{95}/\Gamma_{38}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.193 \pm 0.007</math> OUR FIT</b>				Error includes scale factor of 3.1.
<b><math>0.1901 \pm 0.0024</math> OUR AVERAGE</b>				
$0.1899 \pm 0.0011 \pm 0.0022$	101k $\pm$ 561	WON	09	BELL $e^+e^-$ at $\Upsilon(4S)$
$0.1892 \pm 0.0155 \pm 0.0073$	278 $\pm$ 21	ARMS	04	CLEO $e^+e^-$ $\approx 10$ GeV
$0.1996 \pm 0.0119 \pm 0.0096$	949	LINK	02b	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$0.222 \pm 0.037 \pm 0.013$	63 $\pm$ 10	ABLIKIM	05f	BES $e^+e^- \approx \psi(3770)$
$0.222 \pm 0.041 \pm 0.019$	70	BISHAI	97	CLEO See ARMS 04
$0.25 \pm 0.04 \pm 0.02$	129	FRABETTI	95	E687 $\gamma\text{Be}, \bar{E}_\gamma \approx 200$ GeV
$0.271 \pm 0.065 \pm 0.039$	69	ANJOS	90c	E691 $\gamma\text{Be}$
$0.317 \pm 0.086 \pm 0.048$	31	BALTRUSAIT..85E	MRK3	$e^+e^-$ 3.77 GeV
$0.25 \pm 0.15$	6	SCHINDLER	81	MRK2 $e^+e^-$ 3.771 GeV

 $\Gamma(K^+K_S^0)/\Gamma(K^-2\pi^+)$   $\Gamma_{95}/\Gamma_{40}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.12 \pm 0.16</math> OUR FIT</b>				Error includes scale factor of 3.2.
<b><math>3.35 \pm 0.06 \pm 0.07</math></b>	5161 $\pm$ 86	MENDEZ	10	CLEO $e^+e^-$ at 3774 MeV
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$3.02 \pm 0.18 \pm 0.15$	949	<sup>1</sup> LINK	02b	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

<sup>1</sup> This LINK 02b result is redundant with a result in the previous datablock. $\Gamma(K^+K^-\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{96}/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.996 \pm 0.026</math> OUR FIT</b>				Error includes scale factor of 1.3.
<b><math>0.981 \pm 0.010 \pm 0.032</math></b>		BONVICINI	14	CLEO All CLEO-c runs
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$0.935 \pm 0.017 \pm 0.024$		<sup>1</sup> DOBBS	07	CLEO See BONVICINI 14
$0.97 \pm 0.04 \pm 0.04$	1250 $\pm$ 40	<sup>1</sup> HE	05	CLEO See DOBBS 07

<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05. $\Gamma(K^+K^-\pi^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{96}/\Gamma_{40}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.1053 \pm 0.0024</math> OUR FIT</b>				Error includes scale factor of 1.3.
<b><math>0.1058 \pm 0.0029</math> OUR AVERAGE</b>				Error includes scale factor of 1.4.
$0.117 \pm 0.013 \pm 0.007$	181 $\pm$ 20	ABLIKIM	05f	BES $e^+e^- \approx \psi(3770)$
$0.107 \pm 0.001 \pm 0.002$	43k	AUBERT	05s	BABR $e^+e^- \approx \Upsilon(4S)$
$0.093 \pm 0.010 \pm 0.008$		JUN	00	SELX $\Sigma^-$ nucleus, 600 GeV
$0.0976 \pm 0.0042 \pm 0.0046$		FRABETTI	95b	E687 $\gamma\text{Be}, \bar{E}_\gamma \approx 200$ GeV

 $\Gamma(\phi\pi^+, \phi \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{97}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>27.8 \pm 0.4 \pm 0.2</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	
$29.2 \pm 3.1 \pm 3.0$	FRABETTI	95b	E687 Dalitz fit, 915 evts

 $\Gamma(K^+\bar{K}^*(892)^0, \bar{K}^*(892)^0 \rightarrow K^-\pi^+)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{98}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>25.7 \pm 0.5 \pm 0.4</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	
$30.1 \pm 2.0 \pm 2.5$	FRABETTI	95b	E687 Dalitz fit, 915 evts

 $\Gamma(K^+\bar{K}_S^0(1430)^0, \bar{K}_S^0(1430)^0 \rightarrow K^-\pi^+)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{99}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>18.8 \pm 1.2 \pm 3.3</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	
$37.0 \pm 3.5 \pm 1.8$	FRABETTI	95b	E687 Dalitz fit, 915 evts

 $\Gamma(K^+\bar{K}_S^*(1430)^0, \bar{K}_S^*(1430)^0 \rightarrow K^-\pi^+)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{100}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>1.7 \pm 0.4 \pm 1.2</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts

 $\Gamma(K^+\bar{K}_S^*(800), \bar{K}_S^*(800) \rightarrow K^-\pi^+)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{101}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>7.0 \pm 0.8 \pm 3.5</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts

 $\Gamma(a_0(1450)^0\pi^+, a_0^0 \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{102}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>4.6 \pm 0.6 \pm 7.2</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts

 $\Gamma(\phi(1680)\pi^+, \phi \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{103}/\Gamma_{96}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>0.51 \pm 0.11 \pm 0.37</math></b>	RUBIN	08	CLEO Dalitz fit, 19,458 $\pm$ 163 evts

 $\Gamma(K^*(892) + K_S^0)/\Gamma(K_S^0\pi^+)$   $\Gamma_{111}/\Gamma_{38}$ Unseen decay modes of the  $K^*(892)^+$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 0.3 \pm 0.4</math></b>	67	FRABETTI	95	E687 $\gamma\text{Be}, \bar{E}_\gamma \approx 200$ GeV

 $\Gamma(\phi\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{108}/\Gamma$ Unseen decay modes of the  $\phi$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.023 \pm 0.010</math></b>	<sup>1</sup> BARLAG	92c	ACCM $\pi^-$ Cu 230 GeV

<sup>1</sup> BARLAG 92c computes the branching fraction using topological normalization. $\Gamma(\phi\rho^+)/\Gamma(K^-2\pi^+)$   $\Gamma_{109}/\Gamma_{40}$ Unseen decay modes of the  $\phi$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.16</math></b>	90	DAOUDI	92	CLEO $e^+e^- \approx 10.5$ GeV

 $\Gamma(K^+K^-\pi^+\pi^0\text{non-}\phi)/\Gamma_{\text{total}}$   $\Gamma_{110}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.015 \pm 0.007</math></b>	<sup>1</sup> BARLAG	92c	ACCM $\pi^-$ Cu 230 GeV

<sup>1</sup> BARLAG 92c computes the branching fraction using topological normalization. $\Gamma(K^+K^-\pi^+\pi^0\text{non-}\phi)/\Gamma(K^-2\pi^+)$   $\Gamma_{110}/\Gamma_{40}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$<0.25$	90	ANJOS	89E	E691 Photoproduction

 $\Gamma(K^+K_S^0\pi^+\pi^-)/\Gamma(K_S^02\pi^+\pi^-)$   $\Gamma_{105}/\Gamma_{61}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.62 \pm 0.39 \pm 0.40</math></b>	469 $\pm$ 32	LINK	01c	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

# Meson Particle Listings

$D^\pm$

$\Gamma(K_S^0 K^- 2\pi^+)/\Gamma(K_S^0 2\pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{106}/\Gamma_{61}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.68±0.41±0.32</b>	670 ± 35	LINK	01c FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV

$\Gamma(K^+ K^- 2\pi^+ \pi^-)/\Gamma(K^- 3\pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{107}/\Gamma_{62}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.040±0.009±0.019</b>	38	LINK	03d FOCS	$\gamma$ A, $\overline{E}_\gamma \approx 180$ GeV

Doubly Cabibbo-suppressed modes

$\Gamma(K^+ \pi^0)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{112}/\Gamma</math></span>				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.89±0.25 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>2.52±0.47±0.26</b>	189 ± 37	AUBERT,B	06f BABR	$e^+ e^- \approx \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.28±0.36±0.17	148 ± 23	DYTMAN	06 CLEO	See MENDEZ 10

$\Gamma(K^+ \pi^0)/\Gamma(K^- 2\pi^+)$ <span style="float:right"><math>\Gamma_{112}/\Gamma_{40}</math></span>				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.00±0.26 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>1.9 ±0.2 ±0.1</b>	343 ± 37	MENDEZ	10 CLEO	$e^+ e^-$ at 3774 MeV

$\Gamma(K^+ \eta)/\Gamma(\eta \pi^+)$ <span style="float:right"><math>\Gamma_{113}/\Gamma_{90}</math></span>				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.06±0.43±0.14</b>	166 ± 23	WON	11 BELL	$e^+ e^- \approx \mathcal{T}(4S)$

$\Gamma(K^+ \eta)/\Gamma(K^- 2\pi^+)$ <span style="float:right"><math>\Gamma_{113}/\Gamma_{40}</math></span>				
Unseen decay modes of the $\eta$ are included.				
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.15	90	MENDEZ	10 CLEO	$e^+ e^-$ at 3774 MeV

$\Gamma(K^+ \eta'(958))/\Gamma(\eta'(958) \pi^+)$ <span style="float:right"><math>\Gamma_{114}/\Gamma_{93}</math></span>				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.77±0.39±0.10</b>	180 ± 19	WON	11 BELL	$e^+ e^- \approx \mathcal{T}(4S)$

$\Gamma(K^+ \eta'(958))/\Gamma(K^- 2\pi^+)$ <span style="float:right"><math>\Gamma_{114}/\Gamma_{40}</math></span>				
Unseen decay modes of the $\eta'(958)$ are included.				
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.20	90	MENDEZ	10 CLEO	$e^+ e^-$ at 3774 MeV

$\Gamma(K^+ \pi^+ \pi^-)/\Gamma(K^- 2\pi^+)$ <span style="float:right"><math>\Gamma_{115}/\Gamma_{40}</math></span>				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.77±0.22 OUR AVERAGE</b>				
5.69±0.18±0.14	2638 ± 84	KO	09 BELL	$e^+ e^-$ at $\mathcal{T}(4S)$
6.5 ±0.8 ±0.4	189 ± 24	LINK	04f FOCS	$\gamma$ A, $\overline{E}_\gamma \approx 180$ GeV
7.7 ±1.7 ±0.8	59 ± 13	AITALA	97c E791	$\pi^-$ A, 500 GeV
7.2 ±2.3 ±1.7	21	FRABETTI	95e E687	$\gamma$ Be, $\overline{E}_\gamma = 220$ GeV

$\Gamma(K^+\rho^0)/\Gamma(K^+\pi^+\pi^-)$					$\Gamma_{116}/\Gamma_{115}$
This is the “fit fraction” from the Dalitz-plot analysis.					
VALUE	DOCUMENT ID		TECN	COMMENT	
<b>0.39 ±0.09</b>	<b>OUR AVERAGE</b>				
0.3943 ±0.0787 ±0.0815	LINK	04f	FOCS	Dalitz fit, 189 evts	
0.37 ±0.14 ±0.07	AITALA	97c	E791	Dalitz fit, 59 evts	

$\Gamma(K^+ f_0(980), f_0(980) \rightarrow \pi^+ \pi^-)/\Gamma(K^+ \pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{118}/\Gamma_{115}</math></span>				
This is the “fit fraction” from the Dalitz-plot analysis.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0892±0.0333±0.0412</b>	LINK	04f FOCS	Dalitz fit, 189 evts	

$\Gamma(K^*(892)^0\pi^+, K^*(892)^0 \rightarrow K^+\pi^-)/\Gamma(K^+\pi^+\pi^-)$				$\Gamma_{117}/\Gamma_{115}$
This is the “fit fraction” from the Dalitz-plot analysis.				
VALUE		DOCUMENT ID	TECN	COMMENT
<b>0.47 ±0.08</b>	<b>OUR AVERAGE</b>			
0.5220 ±0.0684 ±0.0638		LINK	04f FOCS	Dalitz fit, 189 evts
0.35 ±0.14 ±0.01		AITALA	97c E791	Dalitz fit, 59 evts

$\Gamma(K_2^*(1430)^0 \pi^+, K_2^*(1430)^0 \rightarrow K^+ \pi^-)/\Gamma(K^+ \pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{119}/\Gamma_{115}</math></span>				
This is the “fit fraction” from the Dalitz-plot analysis.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0803±0.0372±0.0391</b>	LINK	04f FOCS	Dalitz fit, 189 evts	

$\Gamma(K^+\pi^+\pi^- \text{ nonresonant})/\Gamma(K^+\pi^+\pi^-)$				$\Gamma_{120}/\Gamma_{115}$
This is the “fit fraction” from the Dalitz-plot analysis.				
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.36 \pm 0.14 \pm 0.07$	<sup>1</sup> AITALA	97C E791	Dalitz fit, 59 evts	

<sup>1</sup> LINK 04f, with three times as many events, finds no need for a nonresonant amplitude.

$\Gamma(2K^+ K^-)/\Gamma(K^- 2\pi^+)$ <span style="float:right"><math>\Gamma_{121}/\Gamma_{40}</math></span>				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.49±2.17±0.22</b>	65	<sup>1</sup> LINK	02i FOCS	$\gamma$ nucleus, $\approx 180$ GeV

<sup>1</sup> LINK 02i finds little evidence for  $\phi K^+$  or  $f_0(980) K^+$  submodes.

Rare or forbidden modes

$\Gamma(\pi^+ e^+ e^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{122}/\Gamma</math></span>				
A test for the $\Delta C = 1$ weak neutral current. Allowed by higher-order electroweak interactions.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 × 10<sup>-6</sup></b>	90	LEES	11g BABR	$e^+ e^- \approx \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.9 × 10 <sup>-6</sup>	90	<sup>1</sup> RUBIN	10 CLEO	$e^+ e^-$ at $\psi(3770)$
<7.4 × 10 <sup>-6</sup>	90	HE	05A CLEO	See RUBIN 10
<5.2 × 10 <sup>-5</sup>	90	AITALA	99g E791	$\pi^- N$ 500 GeV
<1.1 × 10 <sup>-4</sup>	90	FRABETTI	97b E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
<6.6 × 10 <sup>-5</sup>	90	AITALA	96 E791	$\pi^- N$ 500 GeV
<2.5 × 10 <sup>-3</sup>	90	WEIR	90b MRK2	$e^+ e^-$ 29 GeV
<2.6 × 10 <sup>-3</sup>	90	HAAS	88 CLEO	$e^+ e^-$ 10 GeV

<sup>1</sup> This RUBIN 10 limit is for the  $e^+ e^-$  mass in the continuum away from the  $\phi(1020)$ . See the next data block.

$\Gamma(\pi^+ \phi, \phi \rightarrow e^+ e^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{123}/\Gamma</math></span>				
This is <i>not</i> a test for the $\Delta C = 1$ weak neutral current, but leads to the $\pi^+ e^+ e^-$ final state.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>(1.7±1.4±0.1) × 10<sup>-6</sup></b>	4	<sup>1</sup> RUBIN	10 CLEO	$e^+ e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
(2.7±3.6±1.8) × 10 <sup>-6</sup>	2	HE	05A CLEO	See RUBIN 10

<sup>1</sup> This RUBIN 10 result is consistent with the known  $D^+ \rightarrow \phi \pi^+$  and  $\phi \rightarrow e^+ e^-$  fractions.

$\Gamma(\pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{124}/\Gamma</math></span>				
A test for the $\Delta C = 1$ weak neutral current. Allowed by higher-order electroweak interactions.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.3 × 10<sup>-8</sup></b>	90	AAIJ	13Af LHCB	$p \overline{p}$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<6.5 × 10 <sup>-6</sup>	90	LEES	11g BABR	$e^+ e^- \approx \mathcal{T}(4S)$
<3.9 × 10 <sup>-6</sup>	90	<sup>1</sup> ABAZOV	08d D0	$p \overline{p}, E_{\text{cm}} = 1.96$ TeV
<8.8 × 10 <sup>-6</sup>	90	LINK	03f FOCS	$\gamma$ A, $\overline{E}_\gamma \approx 180$ GeV
<1.5 × 10 <sup>-5</sup>	90	AITALA	99g E791	$\pi^- N$ 500 GeV
<8.9 × 10 <sup>-5</sup>	90	FRABETTI	97b E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
<1.8 × 10 <sup>-5</sup>	90	AITALA	96 E791	$\pi^- N$ 500 GeV
<2.2 × 10 <sup>-4</sup>	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
<5.9 × 10 <sup>-3</sup>	90	WEIR	90b MRK2	$e^+ e^-$ 29 GeV
<2.9 × 10 <sup>-3</sup>	90	HAAS	88 CLEO	$e^+ e^-$ 10 GeV

<sup>1</sup> This ABAZOV 08d limit is for the  $\mu^+ \mu^-$  mass in the continuum away from the  $\phi(1020)$ . See the next data block.

$\Gamma(\pi^+ \phi, \phi \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{125}/\Gamma</math></span>				
This is <i>not</i> a test for the $\Delta C = 1$ weak neutral current, but leads to the $\pi^+ \mu^+ \mu^-$ final state.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>(1.8±0.5±0.6) × 10<sup>-6</sup></b>	<sup>1</sup> ABAZOV	08d D0	$p \overline{p}, E_{\text{cm}} = 1.96$ TeV	

<sup>1</sup> This ABAZOV 08d value is consistent with the known  $D^+ \rightarrow \phi \pi^+$  and  $\phi \rightarrow \mu^+ \mu^-$  fractions.

$\Gamma(\rho^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{126}/\Gamma</math></span>				
A test for the $\Delta C = 1$ weak neutral current. Allowed by higher-order electroweak interactions.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.6 × 10<sup>-4</sup></b>	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

$\Gamma(K^+ e^+ e^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{127}/\Gamma</math></span>				
Both quarks would have to change flavor for this decay to occur.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0 × 10<sup>-6</sup></b>	90	LEES	11g BABR	$e^+ e^- \approx \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0 × 10 <sup>-6</sup>	90	RUBIN	10 CLEO	$e^+ e^-$ at $\psi(3770)$
<6.2 × 10 <sup>-6</sup>	90	HE	05A CLEO	See RUBIN 10
<2.0 × 10 <sup>-4</sup>	90	AITALA	99g E791	$\pi^- N$ 500 GeV
<2.0 × 10 <sup>-4</sup>	90	FRABETTI	97b E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
<4.8 × 10 <sup>-3</sup>	90	WEIR	90b MRK2	$e^+ e^-$ 29 GeV



### $\Gamma(K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{128}/\Gamma$

Both quarks would have to change flavor for this decay to occur.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.3 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<9.2 \times 10^{-6}$	90	LINK	03F FOCS	$\gamma A, \bar{E}_\gamma \approx 180 \text{ GeV}$
$<4.4 \times 10^{-5}$	90	AITALA	99G E791	$\pi^- N 500 \text{ GeV}$
$<9.7 \times 10^{-5}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.2 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^- \text{ emulsion } 600 \text{ GeV}$
$<9.2 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(\pi^+ e^+ \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{129}/\Gamma$

A test of lepton-family-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.9 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.1 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.3 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(\pi^+ e^- \mu^+)/\Gamma_{\text{total}}$ $\Gamma_{130}/\Gamma$

A test of lepton-family-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.6 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.3 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.3 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(K^+ e^+ \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{131}/\Gamma$

A test of lepton-family-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.2 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.3 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.4 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(K^+ e^- \mu^+)/\Gamma_{\text{total}}$ $\Gamma_{132}/\Gamma$

A test of lepton-family-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.8 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.2 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.4 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(\pi^- 2e^+)/\Gamma_{\text{total}}$ $\Gamma_{133}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-6}$	90	RUBIN	10 CLEO	$e^+ e^- \text{ at } \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.9 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
$<3.6 \times 10^{-6}$	90	HE	05A CLEO	See RUBIN 10
$<9.6 \times 10^{-5}$	90	AITALA	99G E791	$\pi^- N 500 \text{ GeV}$
$<1.1 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<4.8 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(\pi^- 2\mu^+)/\Gamma_{\text{total}}$ $\Gamma_{134}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.2 \times 10^{-8}$	90	AAIJ	13AF LHCb	$pp \text{ at } 7 \text{ TeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.0 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
$<4.8 \times 10^{-6}$	90	LINK	03F FOCS	$\gamma A, \bar{E}_\gamma \approx 180 \text{ GeV}$
$<1.7 \times 10^{-5}$	90	AITALA	99G E791	$\pi^- N 500 \text{ GeV}$
$<8.7 \times 10^{-5}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<2.2 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^- \text{ emulsion } 600 \text{ GeV}$
$<6.8 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(\pi^- e^+ \mu^+)/\Gamma_{\text{total}}$ $\Gamma_{135}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.0 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5.0 \times 10^{-5}$	90	AITALA	99G E791	$\pi^- N 500 \text{ GeV}$
$<1.1 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.7 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(\rho^- 2\mu^+)/\Gamma_{\text{total}}$ $\Gamma_{136}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.6 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^- \text{ emulsion } 600 \text{ GeV}$

### $\Gamma(K^- 2e^+)/\Gamma_{\text{total}}$ $\Gamma_{137}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<0.9 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.5 \times 10^{-6}$	90	RUBIN	10 CLEO	$e^+ e^- \text{ at } \psi(3770)$
$<4.5 \times 10^{-6}$	90	HE	05A CLEO	See RUBIN 10
$<1.2 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<9.1 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(K^- 2\mu^+)/\Gamma_{\text{total}}$ $\Gamma_{138}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<10 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.3 \times 10^{-5}$	90	LINK	03F FOCS	$\gamma A, \bar{E}_\gamma \approx 180 \text{ GeV}$
$<1.2 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<3.2 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^- \text{ emulsion } 600 \text{ GeV}$
$<4.3 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(K^- e^+ \mu^+)/\Gamma_{\text{total}}$ $\Gamma_{139}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.9 \times 10^{-6}$	90	LEES	11G BABR	$e^+ e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.3 \times 10^{-4}$	90	FRABETTI	97B E687	$\gamma \text{ Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
$<4.0 \times 10^{-3}$	90	WEIR	90B MRK2	$e^+ e^- 29 \text{ GeV}$

### $\Gamma(K^*(892)^- 2\mu^+)/\Gamma_{\text{total}}$ $\Gamma_{140}/\Gamma$

A test of lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.5 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^- \text{ emulsion } 600 \text{ GeV}$

## $D^\pm$ CP-VIOLATING DECAY-RATE ASYMMETRIES

This is the difference between  $D^+$  and  $D^-$  partial widths for the decay to state  $f$ , divided by the sum of the widths:

$$A_{CP}(f) = [\Gamma(D^+ \rightarrow f) - \Gamma(D^- \rightarrow \bar{f})] / [\Gamma(D^+ \rightarrow f) + \Gamma(D^- \rightarrow \bar{f})].$$

### $A_{CP}(\mu^\pm \nu)$ in $D^+ \rightarrow \mu^+ \nu_\mu$ , $D^- \rightarrow \mu^- \bar{\nu}_\mu$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$+8 \pm 8$	EISENSTEIN 08	CLEO	$e^+ e^- \text{ at } \psi(3770)$

### $A_{CP}(K_L^0 e^\pm \nu)$ in $D^+ \rightarrow K_L^0 e^+ \nu_e$ , $D^- \rightarrow K_L^0 e^- \bar{\nu}_e$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-0.59 \pm 0.60 \pm 1.48$	ABLIKIM 15AF	BES3	$e^+ e^- 3773 \text{ MeV}$

### $A_{CP}(K_S^0 \pi^\pm)$ in $D^\pm \rightarrow K_S^0 \pi^\pm$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.41 \pm 0.09$	OUR AVERAGE			
$-1.1 \pm 0.6 \pm 0.2$		BONVICINI 14	CLEO	All CLEO-c runs
$-0.363 \pm 0.094 \pm 0.067$	1738k	<sup>1</sup> KO 12A	BELL	$e^+ e^- \approx \tau(nS)$
$-0.44 \pm 0.13 \pm 0.10$	807k	DEL-AMO-SA..11H	BABR	$e^+ e^- \approx \tau(4S)$
$-1.6 \pm 1.5 \pm 0.9$	10.6k	<sup>2</sup> LINK 02B	FOCS	$\gamma \text{ nucleus}, \bar{E}_\gamma \approx 180 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.71 \pm 0.19 \pm 0.20$		KO 10	BELL	See KO 12A
$-1.3 \pm 0.7 \pm 0.3$	30k	MENDEZ 10	CLEO	See BONVICINI 14
$-0.6 \pm 1.0 \pm 0.3$		DOBBS 07	CLEO	See MENDEZ 10

<sup>1</sup> KO 12A finds that after subtracting the contribution due to  $K^0 - \bar{K}^0$  mixing, the CP asymmetry due to the change of charm is  $(-0.024 \pm 0.094 \pm 0.067)\%$ , consistent with zero.

<sup>2</sup> LINK 02B measures  $N(D^+ \rightarrow K_S^0 \pi^+)/N(D^+ \rightarrow K^- \pi^+ \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .

### $A_{CP}(K^\mp 2\pi^\pm)$ in $D^+ \rightarrow K^- 2\pi^+$ , $D^- \rightarrow K^+ 2\pi^-$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.18 \pm 0.16$	OUR AVERAGE			
$-0.16 \pm 0.15 \pm 0.09$	2.3M	ABAZOV 14L	D0	$p\bar{p}, \sqrt{s} = 1.96 \text{ TeV}$
$-0.3 \pm 0.2 \pm 0.4$		BONVICINI 14	CLEO	All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.1 \pm 0.4 \pm 0.9$	231k	MENDEZ 10	CLEO	See BONVICINI 14
$-0.5 \pm 0.4 \pm 0.9$		DOBBS 07	CLEO	See MENDEZ 10

### $A_{CP}(K^\mp \pi^\pm \pi^\pm \pi^0)$ in $D^+ \rightarrow K^- \pi^+ \pi^+ \pi^0$ , $D^- \rightarrow K^+ \pi^- \pi^- \pi^0$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-0.3 \pm 0.6 \pm 0.4$	BONVICINI 14	CLEO	All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.0 \pm 0.9 \pm 0.9$	DOBBS 07	CLEO	See BONVICINI 14

### $A_{CP}(K_S^0 \pi^\pm \pi^0)$ in $D^+ \rightarrow K_S^0 \pi^+ \pi^0$ , $D^- \rightarrow K_S^0 \pi^- \pi^0$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-0.1 \pm 0.7 \pm 0.2$	BONVICINI 14	CLEO	All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.3 \pm 0.9 \pm 0.3$	DOBBS 07	CLEO	See BONVICINI 14

# Meson Particle Listings

## $D^\pm$

$A_{CP}(K_S^0 \pi^\pm \pi^\pm \pi^-)$ in $D^+ \rightarrow K_S^0 \pi^+ \pi^+ \pi^-$ , $D^- \rightarrow K_S^0 \pi^- \pi^- \pi^+$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>0.0 \pm 1.2 \pm 0.3</math></b>	BONVICINI	14	CLEO	All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.1 \pm 1.1 \pm 0.6$	DOBBS	07	CLEO	See BONVICINI 14

$A_{CP}(\pi^\pm \pi^0)$ in $D^\pm \rightarrow \pi^\pm \pi^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>+2.9 \pm 2.9 \pm 0.3</math></b>	2.6k	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV

$A_{CP}(\pi^\pm \eta)$ in $D^\pm \rightarrow \pi^\pm \eta$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.0 \pm 1.5</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.			
$+1.74 \pm 1.13 \pm 0.19$		WON	11	BELL $e^+ e^- \approx \mathcal{T}(4S)$
$-2.0 \pm 2.3 \pm 0.3$	2.9k	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV

$A_{CP}(\pi^\pm \eta'(958))$ in $D^\pm \rightarrow \pi^\pm \eta'(958)$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.5 \pm 1.2</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.			
$-0.12 \pm 1.12 \pm 0.17$		WON	11	BELL $e^+ e^- \approx \mathcal{T}(4S)$
$-4.0 \pm 3.4 \pm 0.3$	1.0k	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV

$A_{CP}(\bar{K}^0 / K^0 K^\pm)$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.11 \pm 0.17</math> OUR AVERAGE</b>				
$0.03 \pm 0.17 \pm 0.14$	1.0M	<sup>1</sup> AAIJ	14Bd	LHCB $pp$ at 7, 8 TeV
$0.08 \pm 0.28 \pm 0.14$	277k	KO	13	BELL $e^+ e^-$ at $\mathcal{T}(4S)$
$0.46 \pm 0.36 \pm 0.25$	159k	LEES	13e	BABR $e^+ e^-$ at $\mathcal{T}(4S)$

<sup>1</sup>AAIJ 14Bd reports its result as  $A_{CP}(D^\pm \rightarrow K_S^0 \pi^\pm)$  with  $CP$ -violation effects in the  $K^0 - \bar{K}^0$  system subtracted. It also measures  $A_{CP}(D^\pm \rightarrow \bar{K}^0 / K^0 K^\pm) + A_{CP}(D_S^\pm \rightarrow \bar{K}^0 / K^0 \pi^\pm) = (0.41 \pm 0.49 \pm 0.26)\%$ .

$A_{CP}(K_S^0 K^\pm)$ in $D^\pm \rightarrow K_S^0 K^\pm$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.11 \pm 0.25</math> OUR AVERAGE</b>				
$-0.25 \pm 0.28 \pm 0.14$	277k	KO	13	BELL $e^+ e^-$ at $\mathcal{T}(nS)$
$0.13 \pm 0.36 \pm 0.25$	159k	LEES	13e	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$-0.2 \pm 1.5 \pm 0.9$	5.2k	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV
$7.1 \pm 6.1 \pm 1.2$	949	<sup>1</sup> LINK	02b	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.16 \pm 0.58 \pm 0.25$		KO	10	BELL $e^+ e^- \approx \mathcal{T}(4S)$
$6.9 \pm 6.0 \pm 1.5$	949	<sup>2</sup> LINK	02b	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

<sup>1</sup>LINK 02b measures  $N(D^+ \rightarrow K_S^0 K^+)/N(D^+ \rightarrow K_S^0 \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .  
<sup>2</sup>LINK 02b measures  $N(D^+ \rightarrow K_S^0 K^+)/N(D^+ \rightarrow K^- \pi^+ \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .

$A_{CP}(K^+ K^- \pi^\pm)$ in $D^\pm \rightarrow K^+ K^- \pi^\pm$				
See also AAIJ 11G for a search for $CP$ asymmetry in the $D^\pm \rightarrow K^+ K^- \pi^\pm$ Dalitz plots using 370k decays and four different binning schemes. No evidence for $CP$ asymmetry was found.				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.37 \pm 0.29</math> OUR AVERAGE</b>				
$0.37 \pm 0.30 \pm 0.15$	224k	<sup>1</sup> LEES	13F	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$-0.03 \pm 0.84 \pm 0.29$		RUBIN	08	CLEO $e^+ e^-$ at 3774 MeV
$1.4 \pm 1.0 \pm 0.8$	43k	<sup>2</sup> AUBERT	05s	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$0.6 \pm 1.1 \pm 0.5$	14k	<sup>3</sup> LINK	00b	FOCS
$-1.4 \pm 2.9$		<sup>3</sup> AITALA	97B	E791 $-0.062 < A_{CP} < +0.034$ (90% CL)
$-3.1 \pm 6.8$		<sup>3</sup> FRABETTI	94i	E687 $-0.14 < A_{CP} < +0.081$ (90% CL)
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.1 \pm 0.9 \pm 0.4$		<sup>4</sup> BONVICINI	14	CLEO See RUBIN 08
$-0.1 \pm 1.5 \pm 0.8$		DOBBS	07	CLEO See BONVICINI 14 and RUBIN 08

<sup>1</sup>This is the integrated  $CP$  asymmetry. LEES 13F also searches for  $CP$  asymmetries in four regions of the Dalitz plots (two of which are listed below); in comparisons of binned  $D^+$  and  $D^-$  Dalitz plots; in parametrized fits to those plots, including 2-body submodes; and in comparisons of Legendre-polynomial distributions for the  $K^+ K^-$  and  $K^- \pi^+$  systems.  
<sup>2</sup>AUBERT 05s measures  $N(D^+ \rightarrow K^+ K^- \pi^+)/N(D_S^+ \rightarrow K^+ K^- \pi^+)$ , the ratio of the numbers of events observed, and similarly for the  $D^-$ .  
<sup>3</sup>FRABETTI 94i, AITALA 98c, and LINK 00b measure  $N(D^+ \rightarrow K^- K^+ \pi^+)/N(D^+ \rightarrow K^- \pi^+ \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .  
<sup>4</sup>RUBIN 08 performs a dedicated analysis of this decay mode on the same dataset, with slightly better precision. We therefore take it that BONVICINI 14 does not supersede RUBIN 08's  $A_{CP}$  result.

$A_{CP}(K^\pm K^*0)$ in $D^+ \rightarrow K^+ \bar{K}^{*0}$ , $D^- \rightarrow K^- K^{*0}$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.3 \pm 0.4</math> OUR AVERAGE</b>				
$-0.3 \pm 0.4 \pm 0.2$	73k	<sup>1</sup> LEES	13F	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$-0.4 \pm 2.0 \pm 0.6$		RUBIN	08	CLEO Fit-fraction asymmetry
$+0.9 \pm 1.7 \pm 0.7$	11k	<sup>2</sup> AUBERT	05s	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$-1.0 \pm 5.0$		<sup>3</sup> AITALA	97B	E791 $-0.092 < A_{CP} < +0.072$ (90% CL)
$-12 \pm 13$		<sup>3</sup> FRABETTI	94i	E687 $-0.33 < A_{CP} < +0.094$ (90% CL)

<sup>1</sup>This LEES 13F result is for the  $K^\mp \pi^\pm$  mass-squared between 0.4 and 1.0 GeV<sup>2</sup>, and does not actually separate out the  $K^*$ .  
<sup>2</sup>AUBERT 05s measures  $N(D^+ \rightarrow K^+ \bar{K}^{*0})/N(D_S^+ \rightarrow K^+ K^- \pi^+)$ , the ratio of the numbers of events observed, and similarly for the  $D^-$ .  
<sup>3</sup>FRABETTI 94i and AITALA 97B measure  $N(D^+ \rightarrow K^+ \bar{K}^*(892)^0)/N(D^+ \rightarrow K^- \pi^+ \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .

$A_{CP}(\phi \pi^\pm)$ in $D^\pm \rightarrow \phi \pi^\pm$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.09 \pm 0.19</math> OUR AVERAGE</b>				Error includes scale factor of 1.2.
$-0.04 \pm 0.14 \pm 0.14$	1.58M	AAIJ	13w	LHCB $pp$ at 7 TeV
$-0.3 \pm 0.3 \pm 0.5$	97k	<sup>1</sup> LEES	13F	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$+0.51 \pm 0.28 \pm 0.05$	237k	STARIC	12	BELL Mainly at $\mathcal{T}(4S)$
$-1.8 \pm 1.6 \pm 0.2$		RUBIN	08	CLEO Fit-fraction asymmetry
$+0.2 \pm 1.5 \pm 0.6$	10k	<sup>2</sup> AUBERT	05s	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$-2.8 \pm 3.6$		<sup>3</sup> AITALA	97B	E791 $-0.087 < A_{CP} < +0.031$ (90% CL)
$+6.6 \pm 8.6$		<sup>3</sup> FRABETTI	94i	E687 $-0.075 < A_{CP} < +0.21$ (90% CL)

<sup>1</sup>This LEES 13F result is for the  $K^+ K^-$  mass-squared less than 1.3 GeV<sup>2</sup> and the  $K^\mp \pi^\pm$  mass-squared above 1.0 GeV<sup>2</sup>, and does not actually separate out the  $\phi$ .  
<sup>2</sup>AUBERT 05s measures  $N(D^+ \rightarrow \phi \pi^+)/N(D_S^+ \rightarrow K^+ K^- \pi^+)$ , the ratio of the numbers of events observed, and similarly for the  $D^-$ .  
<sup>3</sup>FRABETTI 94i and AITALA 97B measure  $N(D^+ \rightarrow \phi \pi^+)/N(D^+ \rightarrow K^- \pi^+ \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .

$A_{CP}(K^\pm K_S^*(1430)^0)$ in $D^+ \rightarrow K^+ \bar{K}_S^{*0}(1430)^0$ , $D^- \rightarrow K^- K_S^*(1430)^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>+8 \pm 6 \pm 4</math></b>	RUBIN	08	CLEO	Fit-fraction asymmetry

$A_{CP}(K^\pm K_2^*(1430)^0)$ in $D^+ \rightarrow K^+ \bar{K}_2^{*0}(1430)^0$ , $D^- \rightarrow K^- K_2^*(1430)^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>+43 \pm 19 \pm 18</math></b>	RUBIN	08	CLEO	Fit-fraction asymmetry

$A_{CP}(K^\pm K_S^*(800))$ in $D^+ \rightarrow K^+ \bar{K}_S^{*0}(800)$ , $D^- \rightarrow K^- K_S^*(800)$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-12 \pm 11 \pm 16</math></b>	RUBIN	08	CLEO	Fit-fraction asymmetry

$A_{CP}(a_0(1450)^0 \pi^\pm)$ in $D^\pm \rightarrow a_0(1450)^0 \pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-19 \pm 12 \pm 11</math></b>	RUBIN	08	CLEO	Fit-fraction asymmetry

$A_{CP}(\phi(1680) \pi^\pm)$ in $D^\pm \rightarrow \phi(1680) \pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-9 \pm 22 \pm 14</math></b>	RUBIN	08	CLEO	Fit-fraction asymmetry

$A_{CP}(\pi^+ \pi^- \pi^\pm)$ in $D^\pm \rightarrow \pi^+ \pi^- \pi^\pm$				
See also AAIJ 14c for a search for $CP$ violation in $D^\pm \rightarrow \pi^+ \pi^- \pi^\pm$ Dalitz plots using model-independent binned and unbinned methods. No evidence was found.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-1.7 \pm 4.2</math></b>	<sup>1</sup> AITALA	97B	E791	$-0.086 < A_{CP} < +0.052$ (90% CL)

<sup>1</sup>AITALA 97B measure  $N(D^+ \rightarrow \pi^+ \pi^- \pi^+)/N(D^+ \rightarrow K^- \pi^+ \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $D^-$ .

$A_{CP}(K_S^0 K^\pm \pi^+ \pi^-)$ in $D^\pm \rightarrow K_S^0 K^\pm \pi^+ \pi^-$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-4.2 \pm 6.4 \pm 2.2</math></b>	523 $\pm$ 32	LINK	05e	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$A_{CP}(K^\pm \pi^0)$ in $D^\pm \rightarrow K^\pm \pi^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-3.5 \pm 10.7 \pm 0.9</math></b>	343 $\pm$ 37	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV

$D^\pm$   $\chi^2$  TESTS OF  $CP$ -VIOLATION ( $CPV$ )

We list model-independent searches for local  $CP$  violation in phase-space distributions of multi-body decays.

Most of these searches divide phase space (Dalitz plot for 3-body decays, five-dimensional equivalent for 4-body decays) into bins, and perform a  $\chi^2$  test comparing normalised yields  $N_i$ ,  $\bar{N}_i$  in  $CP$ -conjugate bin pairs  $i$ :  $\chi^2 = \sum_i (N_i - \alpha \bar{N}_i) / \sigma(N_i - \alpha \bar{N}_i)$ . The factor  $\alpha = (\sum_i N_i) / (\sum_i \bar{N}_i)$  removes the dependence on phase-space-integrated rate asymmetries. The result is used to obtain the probability (p-value) to obtain the measured  $\chi^2$  or larger under the assumption of  $CP$  conservation [AUBERT 08AO, BEDIAGA 09]. Alternative methods obtain p-values from other test variables based on unbinned analyses [WILLIAMS 11, AAIJ 14c]. Results can be combined using Fisher's method [MOSTELLER 48].

Local  $CPV$  in  $D^\pm \rightarrow \pi^+ \pi^- \pi^\pm$ 

p-value (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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78.1	3.1M	<sup>1</sup> AAIJ	14c LHCb	$\chi^2$
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<sup>1</sup> AAIJ 14c uses binned and unbinned methods, and finds slightly better sensitivity with the former. We took the first value in the table of results for the binned method.

Local  $CPV$  in  $D^\pm \rightarrow K^+ K^- \pi^\pm$ 

p-value (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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31	OUR EVALUATION			
72	224k	LEES	13f BABR	$\chi^2$
12.7	370k	<sup>1</sup> AAIJ	11g LHCb	$\chi^2$

<sup>1</sup> AAIJ 11g publishes results for several binning schemes. We picked the first value in their table of results.

 $CP$  VIOLATING ASYMMETRIES OF  $P$ -ODD ( $T$ -ODD) MOMENTS $A_{T\text{viol}}(K_S^0 K^\pm \pi^+ \pi^-)$  in  $D^\pm \rightarrow K_S^0 K^\pm \pi^+ \pi^-$ 

$C_T \equiv \bar{p}_{K^+} \cdot (\bar{p}_{\pi^+} \times \bar{p}_{\pi^-})$  is a parity-odd correlation of the  $K^+$ ,  $\pi^+$ , and  $\pi^-$  momenta for the  $D^+$ .  $\bar{C}_T \equiv \bar{p}_{K^-} \cdot (\bar{p}_{\pi^-} \times \bar{p}_{\pi^+})$  is the corresponding quantity for the  $D^-$ . Then

$A_T \equiv [\Gamma(C_T > 0) - \Gamma(C_T < 0)] / [\Gamma(C_T > 0) + \Gamma(C_T < 0)]$ , and  
 $\bar{A}_T \equiv [\Gamma(-\bar{C}_T > 0) - \Gamma(-\bar{C}_T < 0)] / [\Gamma(-\bar{C}_T > 0) + \Gamma(-\bar{C}_T < 0)]$ , and

$A_{T\text{viol}} \equiv \frac{1}{2}(A_T - \bar{A}_T)$ .  $C_T$  and  $\bar{C}_T$  are commonly referred to as  $T$ -odd moments, because they are odd under  $T$  reversal. However, the  $T$ -conjugate process  $K_S^0 K^\pm \pi^+ \pi^- \rightarrow D^\pm$  is not accessible, while the  $P$ -conjugate process is.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$-12.0 \pm 10.0 \pm 4.6$	21.2 $\pm 0.4$ k	LEES	11f BABR	$e^+ e^- \approx \Upsilon(4S)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
23 $\pm 62$ $\pm 22$	523 $\pm 32$	LINK	05f FOCUS	$\gamma A, \bar{E}_\gamma \approx 180$ GeV

 $D^+ \rightarrow (\bar{K}^0/\pi^0/\eta/\omega/\rho^0/\bar{K}^{*0})\ell^+ \nu_\ell$  FORM FACTORS $f_+(0)|V_{cd}|$  in  $D^+ \rightarrow \bar{K}^0 \ell^+ \nu_\ell$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.725 \pm 0.015</math> OUR AVERAGE</b>	Error includes scale factor of 1.7.			
$0.737 \pm 0.006 \pm 0.009$	40k	<sup>1</sup> ABLIKIM	15AF BES3	$K_L e^+ \nu_e$ 3-parameter fit
$0.707 \pm 0.010 \pm 0.009$		<sup>2</sup> BESSON	09 CLEO	$K_S e^+ \nu_e$ 3-parameter fit

<sup>1</sup> ABLIKIM 15AF finds  $0.728 \pm 0.006 \pm 0.011$  for a 2-parameter fit.

<sup>2</sup> BESSON 09 finds  $0.716 \pm 0.007 \pm 0.009$  for a 2-parameter fit.

 $r_1 \equiv a_1/a_0$  in  $D^+ \rightarrow \bar{K}^0 \ell^+ \nu_\ell$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-1.8 \pm 0.4</math> OUR AVERAGE</b>				
$-2.23 \pm 0.42 \pm 0.53$	40k	<sup>1</sup> ABLIKIM	15AF BES3	$K_L e^+ \nu_e$ 3-parameter fit
$-1.66 \pm 0.44 \pm 0.10$		<sup>2</sup> BESSON	09 CLEO	$K_S e^+ \nu_e$ 3-parameter fit

<sup>1</sup> ABLIKIM 15AF finds  $r_1 = -1.91 \pm 0.33 \pm 0.28$  for a 2-parameter fit.

<sup>2</sup> BESSON 09 finds  $r_1 = -2.10 \pm 0.25 \pm 0.08$  for 2-parameter fit.

 $r_2 \equiv a_2/a_0$  in  $D^+ \rightarrow \bar{K}^0 \ell^+ \nu_\ell$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-3 \pm 12</math> OUR AVERAGE</b>	Error includes scale factor of 1.5.			
$+11 \pm 9 \pm 9$	40k	ABLIKIM	15AF BES3	$K_L e^+ \nu_e$ 3-parameter fit
$-14 \pm 11 \pm 1$		BESSON	09 CLEO	$K_S e^+ \nu_e$ 3-parameter fit

 $f_+(0)|V_{cd}|$  in  $D^+ \rightarrow \pi^0 \ell^+ \nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.146 \pm 0.007 \pm 0.002</math></b>	BESSON	09 CLEO	$\pi^0 e^+ \nu_e$ 3-parameter fit

 $r_1 \equiv a_1/a_0$  in  $D^+ \rightarrow \pi^0 \ell^+ \nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-1.37 \pm 0.88 \pm 0.24</math></b>	BESSON	09 CLEO	$\pi^0 e^+ \nu_e$ 3-parameter fit

 $r_2 \equiv a_2/a_0$  in  $D^+ \rightarrow \pi^0 \ell^+ \nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-4 \pm 5 \pm 1</math></b>	BESSON	09 CLEO	$\pi^0 e^+ \nu_e$ 3-parameter fit

 $f_+(0)|V_{cd}|$  in  $D^+ \rightarrow \eta e^+ \nu_e$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.086 \pm 0.006 \pm 0.001</math></b>	YELTON	11 CLEO	z expansion

 $r_1 \equiv a_1/a_0$  in  $D^+ \rightarrow \eta e^+ \nu_e$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-1.83 \pm 2.23 \pm 0.28</math></b>	YELTON	11 CLEO	z expansion

 $r_V \equiv V(0)/A_1(0)$  in  $D^+ \rightarrow \omega e^+ \nu_e$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.24 \pm 0.09 \pm 0.06</math></b>	ABLIKIM	15w BES3	292 fb $^{-1}$ , 3773 MeV

 $r_2 \equiv A_2(0)/A_1(0)$  in  $D^+ \rightarrow \omega e^+ \nu_e$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.06 \pm 0.15 \pm 0.05</math></b>	ABLIKIM	15w BES3	292 fb $^{-1}$ , 3773 MeV

 $r_V \equiv V(0)/A_1(0)$  in  $D^+, D^0 \rightarrow \rho e^+ \nu_e$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.48 \pm 0.15 \pm 0.05</math></b>	<sup>1</sup> DOBBS	13 CLEO	$e^+ e^-$ at $\psi(3770)$

<sup>1</sup> Uses both  $D^+$  and  $D^0$  events. Using PDG 10 values of  $V_{cd}$  and lifetimes, DOBBS 13 gets  $A_1(0) = 0.56 \pm 0.01^{+0.02}_{-0.03}$ ,  $A_2(0) = 0.47 \pm 0.06 \pm 0.04$ , and  $V(0) = 0.84 \pm 0.09^{+0.05}_{-0.06}$ .

 $r_2 \equiv A_2(0)/A_1(0)$  in  $D^+, D^0 \rightarrow \rho e^+ \nu_e$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.83 \pm 0.11 \pm 0.04</math></b>	<sup>1</sup> DOBBS	13 CLEO	$e^+ e^-$ at $\psi(3770)$

<sup>1</sup> Uses both  $D^+$  and  $D^0$  events. Using PDG 10 values of  $V_{cd}$  and lifetimes, DOBBS 13 gets  $A_1(0) = 0.56 \pm 0.01^{+0.02}_{-0.03}$ ,  $A_2(0) = 0.47 \pm 0.06 \pm 0.04$ , and  $V(0) = 0.84 \pm 0.09^{+0.05}_{-0.06}$ .

 $r_V \equiv V(0)/A_1(0)$  in  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$ 

See also BRIERE 10 for  $\bar{K}^* \ell^+ \nu_\ell$  helicity-basis form-factor measurements.

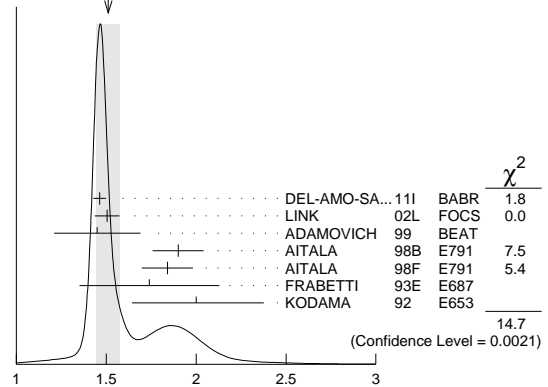
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.51 \pm 0.07</math> OUR AVERAGE</b>	Error includes scale factor of 2.2. See the ideogram below.			
$1.463 \pm 0.017 \pm 0.031$		<sup>1</sup> DEL-AMO-SA...11i	BABR	
$1.504 \pm 0.057 \pm 0.039$	15k	<sup>2</sup> LINK	02L FOCUS	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
$1.45 \pm 0.23 \pm 0.07$	763	ADAMOVICH	99 BEAT	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
$1.90 \pm 0.11 \pm 0.09$	3000	<sup>3</sup> AITALA	98B E791	$\bar{K}^*(892)^0 e^+ \nu_e$
$1.84 \pm 0.11 \pm 0.09$	3034	AITALA	98F E791	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
$1.74 \pm 0.27 \pm 0.28$	874	FRABETTI	93E E687	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
$2.00 \pm 0.34 \pm 0.16$	305	KODAMA	92 E653	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
$2.0 \pm 0.6 \pm 0.3$	183	ANJOS	90E E691	$\bar{K}^*(892)^0 e^+ \nu_e$

<sup>1</sup> DEL-AMO-SANCHEZ 11i finds the pole mass  $m_A = (2.63 \pm 0.10 \pm 0.13)$  GeV ( $m_V$  is fixed at 2 GeV).

<sup>2</sup> LINK 02L includes the effects of interference with an  $S$ -wave background. This much improves the goodness of fit, but does not much shift the values of the form factors.

<sup>3</sup> This is slightly different from the AITALA 98B value: see ref. [5] in AITALA 98F.

WEIGHTED AVERAGE  
 $1.51 \pm 0.07$  (Error scaled by 2.2)



$r_V \equiv V(0)/A_1(0)$  in  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$

 $r_2 \equiv A_2(0)/A_1(0)$  in  $D^+ \rightarrow \bar{K}^*(892)^0 \ell^+ \nu_\ell$ 

See also BRIERE 10 for  $\bar{K}^* \ell^+ \nu_\ell$  helicity-basis form-factor measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.807 \pm 0.025</math> OUR AVERAGE</b>				
$0.801 \pm 0.020 \pm 0.020$		<sup>1</sup> DEL-AMO-SA...11i	BABR	
$0.875 \pm 0.049 \pm 0.064$	15k	<sup>2</sup> LINK	02L FOCUS	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
$1.00 \pm 0.15 \pm 0.03$	763	ADAMOVICH	99 BEAT	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
$0.71 \pm 0.08 \pm 0.09$	3000	AITALA	98B E791	$\bar{K}^*(892)^0 e^+ \nu_e$

# Meson Particle Listings

## D<sup>±</sup>

0.75 ±0.08 ±0.09	3034	AITALA	98F	E791	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
0.78 ±0.18 ±0.10	874	FRABETTI	93E	E687	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
0.82 $^{+0.22}_{-0.23}$ ±0.11	305	KODAMA	92	E653	$\bar{K}^*(892)^0 \mu^+ \nu_\mu$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.0 ±0.5 ±0.2	183	ANJOS	90E	E691	$\bar{K}^*(892)^0 e^+ \nu_e$

- <sup>1</sup> DEL-AMO-SANCHEZ 111 finds the pole mass  $m_A = (2.63 \pm 0.10 \pm 0.13)$  GeV ( $m_V$  is fixed at 2 GeV).
- <sup>2</sup> LINK 02L includes the effects of interference with an S-wave background. This much improves the goodness of fit, but does not much shift the values of the form factors.

### r<sub>3</sub> ≡ A<sub>3</sub>(0)/A<sub>1</sub>(0) in D<sup>+</sup> → K<sup>\*</sup>(892)<sup>0</sup>ℓ<sup>+</sup>ν<sub>ℓ</sub>

See also BRIERE 10 for  $\bar{K}^* \ell^+ \nu_\ell$  helicity-basis form-factor measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.04±0.33±0.29</b>	3034	AITALA	98F	E791 $\bar{K}^*(892)^0 \mu^+ \nu_\mu$

### Γ<sub>L</sub>/Γ<sub>T</sub> in D<sup>+</sup> → K<sup>\*</sup>(892)<sup>0</sup>ℓ<sup>+</sup>ν<sub>ℓ</sub>

See also BRIERE 10 for  $\bar{K}^* \ell^+ \nu_\ell$  helicity-basis form-factor measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.13±0.08 OUR AVERAGE</b>				
1.09±0.10±0.02	763	ADAMOVICH	99	BEAT $\bar{K}^*(892)^0 \mu^+ \nu_\mu$
1.20±0.13±0.13	874	FRABETTI	93E	E687 $\bar{K}^*(892)^0 \mu^+ \nu_\mu$
1.18±0.18±0.08	305	KODAMA	92	E653 $\bar{K}^*(892)^0 \mu^+ \nu_\mu$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.8 $^{+0.6}_{-0.4}$ ±0.3	183	ANJOS	90E	E691 $\bar{K}^*(892)^0 e^+ \nu_e$

### Γ<sub>+</sub>/Γ<sub>-</sub> in D<sup>+</sup> → K<sup>\*</sup>(892)<sup>0</sup>ℓ<sup>+</sup>ν<sub>ℓ</sub>

See also BRIERE 10 for  $\bar{K}^* \ell^+ \nu_\ell$  helicity-basis form-factor measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.22±0.06 OUR AVERAGE Error includes scale factor of 1.6.				
0.28±0.05±0.02	763	ADAMOVICH	99	BEAT $\bar{K}^*(892)^0 \mu^+ \nu_\mu$
0.16±0.05±0.02	305	KODAMA	92	E653 $\bar{K}^*(892)^0 \mu^+ \nu_\mu$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.15 $^{+0.07}_{-0.05}$ ±0.03	183	ANJOS	90E	E691 $\bar{K}^*(892)^0 e^+ \nu_e$

## D<sup>±</sup> REFERENCES

ABLIKIM	15AF	PR D92 112008	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15W	PR D92 071101	M. Ablikim <i>et al.</i>	(BES III Collab.)
AAJ	14BD	JHEP 1410 025	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAJ	14C	PL B728 585	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	14L	PR D90 111102	V. M. Abazov <i>et al.</i>	(DO Collab.)
ABLIKIM	14E	PR D89 052001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	14F	PR D89 051104	M. Ablikim <i>et al.</i>	(BES III Collab.)
BONVICINI	14	PR D89 072002	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
AAJ	13AF	PL B724 203	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAJ	13W	JHEP 1306 112	R. Aaij <i>et al.</i>	(LHCb Collab.)
DOBBS	13	PRL 110 131802	S. Dobbs <i>et al.</i>	(CLEO Collab.)
KO	13	JHEP 1302 098	B.R. Ko <i>et al.</i>	(BELLE Collab.)
LEES	13E	PR D87 052012	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13F	PR D87 052010	J.P. Lees <i>et al.</i>	(BABAR Collab.)
KO	12A	PRL 109 119903 (err.)	B.R. Ko <i>et al.</i>	(BELLE Collab.)
Also		PRL 109 021601	B.R. Ko <i>et al.</i>	(BELLE Collab.)
STARIC	12	PRL 108 071801	M. Staric <i>et al.</i>	(BELLE Collab.)
AAJ	11G	PR D84 112008	R. Aaij <i>et al.</i>	(LHCb Collab.)
DEL-AMO-SA...	11H	PR D83 071103	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	11I	PR D83 072001	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
LEES	11E	PR D84 031103	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	11G	PR D84 072006	J.P. Lees <i>et al.</i>	(BABAR Collab.)
WILLIAMS	11	PR D84 054015	M. Williams	(LOIC)
WON	11	PRL 107 221801	E. Won <i>et al.</i>	(BELLE Collab.)
YELTON	11	PR D84 032001	J. Yelton <i>et al.</i>	(CLEO Collab.)
ANASHIN	10A	PL B686 84	V.V. Anashin <i>et al.</i>	(VEPP-4M)
ASNER	10	PR D81 052007	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	10	PR D81 112001	R.A. Briere <i>et al.</i>	(CLEO Collab.)
KO	10	PRL 104 181602	B.R. Ko <i>et al.</i>	(BELLE Collab.)
MENDEZ	10	PR D81 052013	H. Mendez <i>et al.</i>	(CLEO Collab.)
K PDG	10	JP 037 075021	K. Nakamura <i>et al.</i>	(PDG Collab.)
RUBIN	10	PR D82 092007	P. Rubin <i>et al.</i>	(CLEO Collab.)
BEDIAGA	09	PR D80 096006	I. Bediaga <i>et al.</i>	(CBPF, NDAM)
BESSON	09	PR D80 032005	D. Besson <i>et al.</i>	(CLEO Collab.)
Also		PR D79 052010	J.Y. Ge <i>et al.</i>	(CLEO Collab.)
KO	09	PRL 102 221802	B.R. Ko <i>et al.</i>	(BELLE Collab.)
LINK	09	EPJ B681 14	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
MITCHELL	09B	PRL 102 081801	R.E. Mitchell <i>et al.</i>	(CLEO Collab.)
WON	09	PR D80 111101	E. Won <i>et al.</i>	(BELLE Collab.)
ABAZOV	08D	PRL 100 101801	V.M. Abazov <i>et al.</i>	(DO Collab.)
ABLIKIM	08L	PL B655 16	M. Ablikim <i>et al.</i>	(BES Collab.)
ARTUSO	08	PR D77 092003	M. Artuso <i>et al.</i>	(CLEO Collab.)
AUBERT	08A	PR D78 051102	B. Aubert <i>et al.</i>	(BABAR Collab.)
BONVICINI	08	PR D77 091106	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
BONVICINI	08A	PR D78 052001	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
DOBBS	08	PR D77 112005	S. Dobbs <i>et al.</i>	(CLEO Collab.)
Also		PRL 100 251802	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
EISENSTEIN	08	PR D78 052003	B.I. Eisenstein <i>et al.</i>	(CLEO Collab.)
HE	08	PRL 100 091801	Q. He <i>et al.</i>	(CLEO Collab.)
PDG	08	PL B667 1	C. Amsler <i>et al.</i>	(PDG Collab.)
RUBIN	08	PR D78 072003	P. Rubin <i>et al.</i>	(CLEO Collab.)
ABLIKIM	07	PL B644 20	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	07G	PL B658 1	M. Ablikim <i>et al.</i>	(BES Collab.)
BONVICINI	07	PR D76 012001	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
DOBBS	07	PR D76 112001	S. Dobbs <i>et al.</i>	(CLEO Collab.)
LINK	07B	PL B653 1	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ABLIKIM	06O	EPJ C47 31	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	06P	EPJ C47 39	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	06U	PL B643 246	M. Ablikim <i>et al.</i>	(BES Collab.)
ADAM	06A	PRL 97 251801	N.E. Adam <i>et al.</i>	(CLEO Collab.)
AITALA	06	PR D73 032004	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
Also		PR D74 059901 (err.)	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AUBERT,B	06F	PR D74 011107	B. Aubert <i>et al.</i>	(BABAR Collab.)
DYTMAN	06	PR D74 071102	S.A. Dytman <i>et al.</i>	(CLEO Collab.)
HUANG	06B	PR D74 112005	G.S. Huang <i>et al.</i>	(CLEO Collab.)

LINK	06B	PL B637 32	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
RUBIN	06	PRL 96 081802	P. Rubin <i>et al.</i>	(CLEO Collab.)
RUBIN	06A	PR D73 112005	P. Rubin <i>et al.</i>	(CLEO Collab.)
ABLIKIM	05A	PL B608 24	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05D	PL B610 183	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05F	PL B622 6	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05P	PL B625 196	M. Ablikim <i>et al.</i>	(BES Collab.)
ARTUSO	05A	PRL 95 251801	M. Artuso <i>et al.</i>	(CLEO Collab.)
AUBERT	05S	PRL 95 211101	B. Aubert <i>et al.</i>	(BABAR Collab.)
HE	05	PRL 95 121801	Q. He <i>et al.</i>	(CLEO Collab.)
Also		PRL 96 199903 (err.)	Q. He <i>et al.</i>	(CLEO Collab.)
HE	05A	PRL 95 221802	Q. He <i>et al.</i>	(CLEO Collab.)
HUANG	05B	PRL 95 181801	G.S. Huang <i>et al.</i>	(CLEO Collab.)
KAYIS-TOPAK...	05	PL B626 24	A. Kayis-Topaksu <i>et al.</i>	(CERN CHORUS Collab.)
LINK	05E	PL B622 239	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	05I	PL B621 72	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ABLIKIM	04C	PL B597 39	M. Ablikim <i>et al.</i>	(BEP C BES Collab.)
ARMS	04	PR D69 071102	K. Arms <i>et al.</i>	(CLEO Collab.)
BONVICINI	04A	PR D70 112004	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
LINK	04	PL B585 200	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	04E	PL B598 33	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	04F	PL B601 10	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ANISOVICH	03	EPJ A16 229	V.V. Anisovich <i>et al.</i>	
LINK	03D	PL B561 225	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	03F	PL B572 21	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
AITALA	02	PRL 89 121801	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BRANDENBURG	02	PRL 89 222001	G. Brandenburg <i>et al.</i>	(CLEO Collab.)
KAYIS-TOPAK...	02	PL B549 48	A. Kayis-Topaksu <i>et al.</i>	(CERN CHORUS Collab.)
LINK	02B	PRL 88 041602	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
Also		PRL 88 153903 (err.)	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	02E	PL B535 43	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	02F	PL B537 192	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	02I	PL B541 227	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	02J	PL B541 243	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	02L	PL B544 89	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
AITALA	01B	PRL 86 770	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
LINK	01C	PRL 87 162001	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ABREU	00C	EPJ C12 209	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ASTIER	00D	PL B486 35	P. Astier <i>et al.</i>	(CERN NOMAD Collab.)
JUN	00	PRL 84 1857	S.Y. Jun <i>et al.</i>	(FNAL SELEX Collab.)
LINK	00B	PL B491 232	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
Also		PL B495 443 (err.)	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ABBIENDI	99K	EPJ C8 573	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ADAMOVICH	99	EPJ C6 35	M. Adamovich <i>et al.</i>	(CERN BEATRICE Collab.)
AITALA	99G	PL B462 401	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BONVICINI	99	PRL 82 4586	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
AITALA	98B	PRL 80 1393	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	98C	PL B421 405	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	98F	PL B440 435	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BAI	98B	PL B429 188	J.Z. Bai <i>et al.</i>	(BEP C BES Collab.)
AITALA	97	PL B397 325	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	97B	PL B403 377	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	97C	PL B404 187	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BISHAI	97	PRL 78 3261	M. Bishai <i>et al.</i>	(CLEO Collab.)
FRABETTI	97	PL B391 235	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	97B	PL B398 239	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	97C	PL B401 131	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	97D	PL B407 79	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
AITALA	96	PRL 76 364	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
FRABETTI	95	PL B346 199	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	95B	PL B351 591	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	95E	PL B359 403	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	95	PL B345 85	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ALBRECHT	94I	ZPHY C64 375	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BALEST	94	PRL 72 2328	R. Balest <i>et al.</i>	(CLEO Collab.)
FRABETTI	94D	PL B323 459	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	94G	PL B331 217	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	94I	PR D50 R2953	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
AKERIB	93	PRL 71 3070	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
ANJOS	93	PR D48 56	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
FRABETTI	93E	PL B307 262	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALBRECHT	92F	PL B278 202	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	92C	PR D46 1941	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BARLAG	92C	ZPHY C55 383	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
Also		ZPHY C48 29	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
COFFMAN	92B	PR D45 2196	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
DAOUDI	92	PR D45 3965	M. Daoudi <i>et al.</i>	(CLEO Collab.)
KODAMA	92	PL B274 246	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
KODAMA	92C	PL B286 187	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ADAMOVICH	91	PL B268 142	M.I. Adamovich <i>et al.</i>	(WA82 Collab.)
ALBRECHT	91	PL B255 634	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALVAREZ	91B	ZPHY C50 11	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
AMMAR	91	PR D44 3383	R. Ammar <i>et al.</i>	(CLEO Collab.)
BAI	91	PRL 66 1011	Z. Bai <i>et al.</i>	(Mark III Collab.)
COFFMAN	91	PL B263 135	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
FRABETTI	91	PL B263 584	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALVAREZ	90	ZPHY C47 539	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
ANJOS	90C	PR D41 2705	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ANJOS	90D	PR D42 2414	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ANJOS	90E	PRL 65 2630	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BARLAG	90C	ZPHY C46 563	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
WEIR	90B	PR D41 1384	A.J. Weir <i>et al.</i>	(Mark II Collab.)
ANJOS	89	PRL 62 125	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ANJOS	89B	PRL 62 722	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ANJOS	89E	PL B223 267	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ADLER	88C	PRL 60 89	J. Adler <i>et al.</i>	(Mark III Collab.)
ALBRECHT	88I	PRL B210 267	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
HAAS	88	PRL 60 1614	P. Haas <i>et al.</i>	(CLEO Collab.)
ONG	88	PRL 60 2587	R.A. Ong <i>et al.</i>	(Mark II Collab.)
RAAB	88	PR D37 2391	J.R. Raab <i>et al.</i>	(FNAL E691 Collab.)
ADAMOVICH	87	EPL 4 887	M.I. Adamovich <i>et al.</i>	(Bhonne Emulsion Collab.)
ADLER	87	PL B196 107	J. Adler <i>et al.</i>	(Mark III Collab.)
BARTEL	87	ZPHY C33 339	W. Bartel <i>et al.</i>	(JADE Collab.)
BALTUSAITIS	86E	PRL 56 2140	R.M. Baltusaitis <i>et al.</i>	(Mark III Collab.)
BALTUSAITIS	85B	PRL 54 1976	R.M. Baltusaitis <i>et al.</i>	(Mark III Collab.)
BALTUSAITIS	85E	PRL 55 150	R.M. Baltusaitis <i>et al.</i>	(Mark III Collab.)
ANJOS	87	PL B138 297	W. Bartel <i>et al.</i>	(JADE Collab.)
ADAMOVICH	84	PL 140B 119	M.I. Adamovich <i>et al.</i>	(CERN WA88 Collab.)
ALTHOFF	84G	ZPHY C22 219	M. Althoff <i>et al.</i>	(TASSO Collab.)
DERRICK	84	PRL 53 1971	M. Derrick <i>et al.</i>	(HRS Collab.)
SCHINDLER	81	PR D24 78	R.H. Schindler <i>et al.</i>	(Mark II Collab.)
TRILLING	81	PRPL 75 57	G.H. Trilling <i>et al.</i>	(LBL, UCB I Collab.)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO Collab.)
Also		SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO Collab.)
Also		Translated from YAF 34 147		
GOLDBABER	77	PL B98 503	G. Goldhaber <i>et al.</i>	(Mark I Collab.)
PERUZZI	77	PRL 39 1301	I. Peruzzi <i>et al.</i>	(LGW Collab.)
PICCOLO	77	PL 70B 260	M. Piccolo <i>et al.</i>	(Mark I Collab.)
PERUZZI	76	PRL 37 569	I. Peruzzi <i>et al.</i>	(Mark I Collab.)
MOSTELLER	48	Am.Stat. 3 No.5 30	R.A. Fisher, F. Mosteller	

## OTHER RELATED PAPERS

RICHMAN ROSNER	95	RMP 67 893 CNPP 21 369	J.D. Richman, P.R. Burchat J. Rosner	(UCSB, STAN) (CHIC)
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$$I(J^P) = \frac{1}{2}(0^-)$$

 $D^0$  MASS

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*0}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

Given the recent addition of much more precise measurements, we have omitted all those masses published up through 1990. See any Review before 2015 for those earlier results.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1864.83 ± 0.05 OUR FIT</b>				
<b>1864.84 ± 0.05 OUR AVERAGE</b>				
1864.845 ± 0.025 ± 0.057	63k	<sup>1</sup> TOMARADZE 14		$D^0 \rightarrow K^- 2\pi^+ \pi^-$
1864.75 ± 0.15 ± 0.11		AAIJ 13v	LHCB	$D^0 \rightarrow K^+ 2K^- \pi^+$
1864.841 ± 0.048 ± 0.063	4.3k	<sup>2</sup> LEES 13s	BABR	$e^+ e^-$ at $\Upsilon(4S)$
1865.30 ± 0.33 ± 0.23	0.1k	ANASHIN 10A	KEDR	$e^+ e^-$ at $\psi(3770)$
1864.847 ± 0.150 ± 0.095	0.3k	CAWLFIELD 07	CLEO	$D^0 \rightarrow K_S^0 \phi$

<sup>1</sup> Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration. The largest source of error in the TOMARADZE 14 value is from the uncertainties in the  $K^-$  and  $K_S^0$  masses. The systematic error given above is the addition in quadrature of  $\pm 0.022 \pm 0.053$  MeV, where the second error is from those mass uncertainties.

<sup>2</sup> The largest source of error in the LEES 13s value is from the uncertainty of the  $K^+$  mass. The quoted systematic error is in fact  $\pm 0.043 + 3 (m_{K^+} - 493.677)$ , in MeV.

 $m_{D^\pm} - m_{D^0}$ 

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*0}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4.75 ± 0.08 OUR FIT</b>			
<b>4.76 ± 0.12 ± 0.07</b>	AAIJ 13v	LHCB	$D^+ \rightarrow K^+ K^- \pi^+$

 $D^0$  MEAN LIFE

Measurements with an error  $> 10 \times 10^{-15}$  s have been omitted from the average.

VALUE ( $10^{-15}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>410.1 ± 1.5 OUR AVERAGE</b>				
409.6 ± 1.1 ± 1.5	210k	LINK 02f	FOCS	$\gamma$ nucleus, $\approx 180$ GeV
407.9 ± 6.0 ± 4.3	10k	KUSHNIR... 01	SELX	$K^- \pi^+, K^- \pi^+ \pi^+ \pi^-$
413 ± 3 ± 4	35k	AITALA 99e	E791	$K^- \pi^+$
408.5 ± 4.1 ± 3.5 3.4	25k	BONVICINI 99	CLE2	$e^+ e^- \approx \Upsilon(4S)$
413 ± 4 ± 3	16k	FRABETTI 94d	E687	$K^- \pi^+, K^- \pi^+ \pi^+ \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
424 ± 11 ± 7	5118	FRABETTI 91	E687	$K^- \pi^+, K^- \pi^+ \pi^+ \pi^-$
417 ± 18 ± 15	890	ALVAREZ 90	NA14	$K^- \pi^+, K^- \pi^+ \pi^+ \pi^-$
388 ± 23 -21	641	<sup>1</sup> BARLAG 90c	ACCM	$\pi^- \text{Cu } 230 \text{ GeV}$
480 ± 40 ± 30	776	ALBRECHT 88i	ARG	$e^+ e^- 10 \text{ GeV}$
422 ± 8 ± 10	4212	RAAB 88	E691	Photoproduction
420 ± 50	90	BARLAG 87b	ACCM	$K^-$ and $\pi^- 200 \text{ GeV}$

<sup>1</sup> BARLAG 90c estimate systematic error to be negligible.

 $D^0 - \bar{D}^0$  MIXING

Revised August 2015 by D. M. Asner (Pacific Northwest National Laboratory)

The detailed formalism for  $D^0 - \bar{D}^0$  mixing is presented in the note on “ $CP$  Violation in Meson Decays” in this *Review*. For completeness, we present an overview here. The time evolution of the  $D^0 - \bar{D}^0$  system is described by the Schrödinger equation

$$i \frac{\partial}{\partial t} \begin{pmatrix} D^0(t) \\ \bar{D}^0(t) \end{pmatrix} = \left( \mathbf{M} - \frac{i}{2} \mathbf{\Gamma} \right) \begin{pmatrix} D^0(t) \\ \bar{D}^0(t) \end{pmatrix}, \quad (1)$$

where the  $\mathbf{M}$  and  $\mathbf{\Gamma}$  matrices are Hermitian, and  $CPT$  invariance requires that  $M_{11} = M_{22} \equiv M$  and  $\Gamma_{11} = \Gamma_{22} \equiv \Gamma$ . The

off-diagonal elements of these matrices describe the dispersive and absorptive parts of the mixing.

Because  $CP$  violation is expected to be quite small here, it is convenient to label the mass eigenstates by the  $CP$  quantum number in the limit of  $CP$  conservation. Thus, we write

$$|D_{1,2}\rangle = p|D^0\rangle \pm q|\bar{D}^0\rangle, \quad (2)$$

where

$$\left( \frac{q}{p} \right)^2 = \frac{M_{12}^* - \frac{i}{2}\Gamma_{12}^*}{M_{12} - \frac{i}{2}\Gamma_{12}}. \quad (3)$$

The normalization condition is  $|p|^2 + |q|^2 = 1$ . Our phase convention is  $CP|D^0\rangle = +|\bar{D}^0\rangle$ , and the sign is chosen so that  $D_1$  has  $CP$  even, or nearly so.

The corresponding eigenvalues are

$$\omega_{1,2} \equiv m_{1,2} - \frac{i}{2}\Gamma_{1,2} = \left( M - \frac{i}{2}\Gamma \right) \pm \frac{q}{p} \left( M_{12} - \frac{i}{2}\Gamma_{12} \right), \quad (4)$$

where  $m_{1,2}$  and  $\Gamma_{1,2}$  are the masses and widths of the  $D_{1,2}$ .

We define dimensionless mixing parameters  $x$  and  $y$  by

$$x \equiv (m_1 - m_2)/\Gamma = \Delta m/\Gamma \quad (5)$$

and

$$y \equiv (\Gamma_1 - \Gamma_2)/2\Gamma = \Delta\Gamma/2\Gamma, \quad (6)$$

where  $\Gamma \equiv (\Gamma_1 + \Gamma_2)/2$ . If  $CP$  is conserved, then  $M_{12}$  and  $\Gamma_{12}$  are real,  $\Delta m = 2M_{12}$ ,  $\Delta\Gamma = 2\Gamma_{12}$ , and  $p = q = 1/\sqrt{2}$ . The signs of  $\Delta m$  and  $\Delta\Gamma$  are to be determined experimentally.

The parameters  $x$  and  $y$  are measured in several ways. The most precise values are obtained using the time dependence of  $D$  decays. Since  $D^0 - \bar{D}^0$  mixing is a small effect, the identifying tag of the initial particle as a  $D^0$  or a  $\bar{D}^0$  must be extremely accurate. The usual tag is the charge of the distinctive slow pion in the decay sequence  $D^{*+} \rightarrow D^0 \pi^+$  or  $D^{*-} \rightarrow \bar{D}^0 \pi^-$ . In current experiments, the probability of mistagging is about 0.1%. The large data samples produced at the  $B$ -factories allow the production flavor to also be determined by fully reconstructing charm on the “other side” of the event—significantly reducing the mistag rate [1]. Another tag of comparable accuracy is identification of one of the  $D$ ’s produced from  $\psi(3770) \rightarrow D^0 \bar{D}^0$  decays. Although time-dependent analyses are not possible at symmetric charm-threshold facilities (the  $D^0$  and  $\bar{D}^0$  do not travel far enough), the quantum-coherent  $C = -1$   $\psi(3770) \rightarrow D^0 \bar{D}^0$  state provides time-integrated sensitivity [2,3].

**Time-Dependent Analyses:** We extend the formalism of this *Review*’s note on “ $CP$  Violation in Meson Decays.” In addition to the “right-sign” instantaneous decay amplitudes  $\bar{A}_f \equiv \langle f|H|\bar{D}^0\rangle$  and  $A_{\bar{f}} \equiv \langle \bar{f}|H|D^0\rangle$  for final states  $f = K^+ \pi^-, \dots$  and their  $CP$  conjugate  $\bar{f} = K^- \pi^+, \dots$ , we include “wrong-sign” amplitudes  $\bar{A}_{\bar{f}} \equiv \langle \bar{f}|H|\bar{D}^0\rangle$  and  $A_f \equiv \langle f|H|D^0\rangle$ .

It is conventional to normalize the wrong-sign decay distributions to the integrated rate of right-sign decays and to express time in units of the precisely measured neutral  $D$ -meson mean lifetime,  $\tau_{D^0} = 1/\Gamma = 2/(\Gamma_1 + \Gamma_2)$ . Starting from a pure  $|D^0\rangle$  or  $|\bar{D}^0\rangle$  state at  $t = 0$ , the time-dependent rates of decay

# Meson Particle Listings

## $D^0$

to wrong-sign final states relative to the integrated right-sign decay rates are, to leading order:

$$r(t) \equiv \frac{|\langle f | H | D^0(t) \rangle|^2}{|\bar{A}_f|^2} = \left| \frac{q}{p} \right|^2 \left| g_+(t) \lambda_f^{-1} + g_-(t) \right|^2, \quad (7)$$

and

$$\bar{r}(t) \equiv \frac{|\langle \bar{f} | H | \bar{D}^0(t) \rangle|^2}{|A_{\bar{f}}|^2} = \left| \frac{p}{q} \right|^2 \left| g_+(t) \lambda_{\bar{f}} + g_-(t) \right|^2. \quad (8)$$

where

$$\lambda_f \equiv q\bar{A}_f/pA_f, \quad \lambda_{\bar{f}} \equiv q\bar{A}_{\bar{f}}/pA_{\bar{f}}, \quad (9)$$

and

$$g_{\pm}(t) = \frac{1}{2} (e^{-iz_1 t} \pm e^{-iz_2 t}), \quad z_{1,2} = \frac{\omega_{1,2}}{\Gamma}. \quad (10)$$

Note that a change in the convention for the relative phase of  $D^0$  and  $\bar{D}^0$  would cancel between  $q/p$  and  $\bar{A}_f/A_f$  and leave  $\lambda_f$  unchanged. We expand  $r(t)$  and  $\bar{r}(t)$  to second order in  $x$  and  $y$  for modes in which the ratio of decay amplitudes,  $R_D = |A_f/\bar{A}_f|^2$ , is very small.

**Semileptonic decays:** Consider the final state  $f = K^+ \ell^- \bar{\nu}_\ell$ , where  $A_f = \bar{A}_{\bar{f}} = 0$  in the Standard Model. The final state  $\bar{f}$  is only accessible through mixing and  $r(t)$  is

$$r(t) = |g_-(t)|^2 \left| \frac{q}{p} \right|^2 \approx \frac{e^{-t}}{4} (x^2 + y^2) t^2 \left| \frac{q}{p} \right|^2. \quad (11)$$

For  $\bar{r}(t)$   $q/p$  is replaced by  $p/q$ . In the Standard Model,  $CP$  violation in charm mixing is small and  $|q/p| \approx 1$ . In the limit of  $CP$  conservation,  $r(t) = \bar{r}(t)$ , and the time-integrated mixing rate relative to the time-integrated right-sign decay rate for semileptonic decays is

$$R_M = \int_0^\infty r(t) dt = \left| \frac{q}{p} \right|^2 \frac{x^2 + y^2}{2 + x^2 - y^2} \approx \frac{1}{2} (x^2 + y^2). \quad (12)$$

**Table 1:** Results for  $R_M$  in  $D^0$  semileptonic decays.

Year	Exper.	Final stat.	$R_M (\times 10^{-3})$	90% C.L.
2008	Belle [4]	$K^{(*)+} e^- \bar{\nu}_e$	$0.13 \pm 0.22 \pm 0.20$	$< 0.61 \times 10^{-3}$
2007	BaBar [1]	$K^{(*)+} e^- \bar{\nu}_e$	$0.04^{+0.70}_{-0.60}$	$(-1.3, 1.2) \times 10^{-3}$
2005*	Belle [5]	$K^{(*)+} e^- \bar{\nu}_e$	$0.02 \pm 0.47 \pm 0.14$	$< 1.0 \times 10^{-3}$
2005	CLEO [6]	$K^{(*)+} e^- \bar{\nu}_e$	$1.6 \pm 2.9 \pm 2.9$	$< 7.8 \times 10^{-3}$
2004*	BaBar [7]	$K^{(*)+} e^- \bar{\nu}_e$	$2.3 \pm 1.2 \pm 0.4$	$< 4.2 \times 10^{-3}$
2002*	FOCUS [8]	$K^+ \mu^- \bar{\nu}_\mu$	$-0.76^{+0.99}_{-0.93}$	$< 1.01 \times 10^{-3}$
1996	E791 [9]	$K^+ \ell^- \bar{\nu}_\ell$	$(1.1^{+3.0}_{-2.7}) \times 10^{-3}$	$< 5.0 \times 10^{-3}$
HFAG [10]			$0.13 \pm 0.27$	

\*These measurements are excluded from the HFAG average. The FOCUS result is unpublished, the statistical correlation of the BaBar result with Ref. 1 has not been established, and the Belle result is superseded by Ref. 4. The HFAG average assumes reported statistical and systematic uncertainties are uncorrelated.

Table 1 summarizes results for  $R_M$  from semileptonic decays; the world average from the Heavy Flavor Averaging Group (HFAG) [10] is  $R_M = (1.30 \pm 2.69) \times 10^{-4}$ .

**Wrong-sign decays to hadronic non- $CP$  eigenstates:** Consider the final state  $f = K^+ \pi^-$ , where  $A_f$  is doubly Cabibbo-suppressed. The ratio of decay amplitudes is

$$\frac{A_f}{\bar{A}_f} = -\sqrt{R_D} e^{-i\delta_f}, \quad \left| \frac{A_f}{\bar{A}_f} \right| \sim O(\tan^2 \theta_c), \quad (13)$$

where  $R_D$  is the doubly Cabibbo-suppressed (DCS) decay rate relative to the Cabibbo-favored (CF) rate,  $\delta_f$  is the strong phase difference between DCS and CF processes, and  $\theta_c$  is the Cabibbo angle. The minus sign originates from the sign of  $V_{us}$  relative to  $V_{cd}$ .

We characterize the violation of  $CP$  with the real-valued parameters  $A_M$ ,  $A_D$ , and  $\phi$ . We adopt the parametrization (see Refs. 11 and 12)

$$\left| \frac{q}{p} \right|^2 = \sqrt{\frac{1+A_M}{1-A_M}}, \quad (14)$$

$$\lambda_f^{-1} \equiv \frac{pA_f}{q\bar{A}_f} = -\sqrt{R_D} \left( \frac{(1+A_D)(1-A_M)}{(1-A_D)(1+A_M)} \right)^{1/4} e^{-i(\delta_f+\phi)}, \quad (15)$$

$$\lambda_{\bar{f}} \equiv \frac{q\bar{A}_{\bar{f}}}{pA_{\bar{f}}} = -\sqrt{R_D} \left( \frac{(1-A_D)(1+A_M)}{(1+A_D)(1-A_M)} \right)^{1/4} e^{-i(\delta_f-\phi)}, \quad (16)$$

and  $A_D$  is a measure of direct  $CP$  violation, while  $A_M$  is a measure of  $CP$  violation in mixing. From these relations, we obtain

$$\sqrt{\frac{1+A_D}{1-A_D}} = \frac{|A_f/\bar{A}_f|}{|\bar{A}_{\bar{f}}/A_{\bar{f}}|}, \quad (17)$$

The angle  $\phi$  measures  $CP$  violation in interference between mixing and decay. While  $A_M$  is independent of the decay process,  $A_D$  and  $\phi$ , in general, depend on  $f$ .

In general,  $\lambda_{\bar{f}}$  and  $\lambda_f^{-1}$  are independent complex numbers. More detail on  $CP$  violation in meson decays can be found in Ref. 13. To leading order, for  $A_D$  and  $A_M \ll 1$ ,

$$r(t) = e^{-t} \left[ R_D(1+A_D) + \sqrt{R_D(1+A_M)(1+A_D)} y'_- t + \frac{1}{2} (1+A_M) R_M t^2 \right] \quad (18)$$

and

$$\bar{r}(t) = e^{-t} \left[ R_D(1-A_D) + \sqrt{R_D(1-A_M)(1-A_D)} y'_+ t + \frac{1}{2} (1-A_M) R_M t^2 \right] \quad (19)$$

Here

$$\begin{aligned} y'_\pm &\equiv y' \cos \phi \pm x' \sin \phi \\ &= y \cos(\delta_{K\pi} \mp \phi) - x \sin(\delta_{K\pi} \mp \phi), \end{aligned} \quad (20)$$

where

$$\begin{aligned} x' &\equiv x \cos \delta_{K\pi} + y \sin \delta_{K\pi}, \\ y' &\equiv y \cos \delta_{K\pi} - x \sin \delta_{K\pi}, \end{aligned} \quad (21)$$

and  $R_M = (x^2 + y^2)/2 = (x'^2 + y'^2)/2$  is the mixing rate relative to the time-integrated Cabibbo-favored rate.

The three terms in Eq. (18) and Eq. (19) probe the three fundamental types of  $CP$  violation. In the limit of  $CP$  conservation,  $A_M$ ,  $A_D$ , and  $\phi$  are all zero. Then

$$r(t) = \bar{r}(t) = e^{-t} \left( R_D + \sqrt{R_D} y' t + \frac{1}{2} R_M t^2 \right), \quad (22)$$

and the time-integrated wrong-sign rate relative to the integrated right-sign rate is

$$R = \int_0^\infty r(t) dt = R_D + \sqrt{R_D} y' + R_M. \quad (23)$$

The ratio  $R$  is the most readily accessible experimental quantity. In Table 2 are reported the measurements of  $R$ ,  $R_D$  and  $A_D$  in  $D^0 \rightarrow K^+ \pi^-$ , and their HFAG average [24] from a general fit; that allows for both mixing and  $CP$  violation. Typically, the fit parameters are  $R_D$ ,  $x'^2$ , and  $y'$ . Table 3 summarizes the results for  $x'^2$  and  $y'$ . Allowing for  $CP$  violation, the separate contributions to  $R$  can be extracted by fitting the  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^- \pi^+$  decay rates.

**Table 2:** Results for  $R$ ,  $R_D$ , and  $A_D$  in  $D^0 \rightarrow K^+ \pi^-$ .

Year	Experiment	$R(\times 10^{-3})$	$R_D(\times 10^{-3})$	$A_D(\%)$
2014	Belle [14]	$3.86 \pm 0.06$	$3.53 \pm 0.13$	—
2013	LHCb [15]	—	$3.57 \pm 0.07$	$-0.7 \pm 1.9$
2013	CDF [16]	$4.30 \pm 0.05$	$3.51 \pm 0.35$	—
2012*	LHCb [17]	$4.25 \pm 0.04$	$3.52 \pm 0.15$	—
2007*	CDF [18]	$4.15 \pm 0.10$	$3.04 \pm 0.55$	—
2007	BaBar [19]	$3.53 \pm 0.08 \pm 0.04$	$3.03 \pm 0.16 \pm 0.10$	$-2.1 \pm 5.2 \pm 1.5$
2006*	Belle [20]	$3.77 \pm 0.08 \pm 0.05$	$3.64 \pm 0.17$	$2.3 \pm 4.7$
2005†	FOCUS [21]	$4.29^{+0.63}_{-0.61} \pm 0.28$	$5.17^{+1.47}_{-1.58} \pm 0.76$	$13^{+33}_{-25} \pm 10$
2000†	CLEO [22]	$3.32^{+0.63}_{-0.65} \pm 0.40$	$4.8 \pm 1.2 \pm 0.4$	$-1^{+16}_{-17} \pm 1$
1998†	E791 [23]	$6.8^{+3.4}_{-3.3} \pm 0.7$	—	—
Average			$3.49 \pm 0.04$ [24]	$-0.39^{+1.01}_{-1.05}$ [24]

\*These measurements are excluded from the HFAG average of  $R_D$ . The CDF result is superseded by Ref. 16 and the LHCb is superseded by Ref. 15. The LHCb result is included in the average of  $R$ . The Belle result for  $R$  and  $R_D$  is superseded by Ref. 14.

†These measurements are excluded from the HFAG average due to poor precision.

Extraction of the mixing parameters  $x$  and  $y$  from the results in Table 3 requires knowledge of the relative strong phase  $\delta_{K\pi}$ . An interference effect that provides useful sensitivity to  $\delta_{K\pi}$  arises in the decay chain  $\psi(3770) \rightarrow D^0 \bar{D}^0 \rightarrow (f_{CP})(K^+ \pi^-)$ , where  $f_{CP}$  denotes a  $CP$ -even or -odd eigenstate from  $D^0$  decay, such as  $K^+ K^-$  or  $K_S^0 \pi^0$ , respectively [26]. Here, the amplitude relation

**Table 3:** Results on the time-dependence of  $r(t)$  in  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^- \pi^+$  decays. The Belle 2014, LHCb and CDF results assume no  $CP$  violation. The FOCUS, CLEO, and Belle 2006 results restrict  $x'^2$  to the physical region. The confidence intervals from FOCUS, CLEO, and BaBar are obtained from the fit, whereas Belle uses a Feldman-Cousins method, and CDF uses a Bayesian method.

Year	Exper.	$y'$ (%)	$x'^2 (\times 10^{-3})$
2014*†	Belle [14]	$0.46 \pm 0.34$	$0.09 \pm 0.22$
2013	LHCb [15]	$0.48 \pm 0.10$	$0.055 \pm 0.049$
2013	CDF [16]	$0.43 \pm 0.43$	$0.08 \pm 0.18$
2012*	LHCb [17]	$0.72 \pm 0.24$	$-0.09 \pm 0.13$
2007*	CDF [18]	$0.85 \pm 0.76$	$-0.12 \pm 0.35$
2007	BaBar [19]	$0.97 \pm 0.44 \pm 0.31$	$-0.22 \pm 0.30 \pm 0.21$
2006†	Belle [20]	$-2.8 < y' < 2.1$	$< 0.72$ (95% C.L.)
2005*	FOCUS [21]	$-11.2 < y' < 6.7$	$< 8.0$ (95% C.L.)
2000*	CLEO [22]	$-5.8 < y' < 1.0$	$< 0.81$ (95% C.L.)

\*These measurements are excluded from the HFAG average. The CDF result is superseded by Ref. 16 and the LHCb result has been superseded by Ref. 15. The CLEO and FOCUS results are excluded due to poor precision.

† This Belle result allows for  $CP$  violation. HFAG uses this result for the  $CP$ -violation allowed fit. This result is not superseded by Ref. 14.

\*† This Belle result does not allow for  $CP$  violation. HFAG uses this result for the  $CP$ -conserving fit. This result does not supersede Ref. 20.

$$\sqrt{2} A(D_\pm \rightarrow K^- \pi^+) = A(D^0 \rightarrow K^- \pi^+) \pm A(\bar{D}^0 \rightarrow K^- \pi^+). \quad (24)$$

where  $D_\pm$  denotes a  $CP$ -even or -odd eigenstate, implies that

$$\cos \delta_{K\pi} = \frac{|A(D_+ \rightarrow K^- \pi^+)|^2 - |A(D_- \rightarrow K^- \pi^+)|^2}{2\sqrt{R_D} |A(D^0 \rightarrow K^- \pi^+)|^2}. \quad (25)$$

This neglects  $CP$  violation and uses  $\sqrt{R_D} \ll 1$ .

The asymmetry of  $CP$ -tagged  $D$  decays rates to  $K^- \pi^+$  is denoted as

$$A_{K\pi}^{CP} \equiv \frac{|A(D_- \rightarrow K^- \pi^+)|^2 - |A(D_+ \rightarrow K^- \pi^+)|^2}{|A(D_- \rightarrow K^- \pi^+)|^2 + |A(D_+ \rightarrow K^- \pi^+)|^2}. \quad (26)$$

To lowest order in the mixing parameters [2,3]

$$2\sqrt{R_D} \cos \delta_{K\pi} + y = (1 + R) \dot{A}_{K\pi}^{CP} \quad (27)$$

where  $R$  is the time-integrated wrong-sign rate relative to the integrated right-sign rate from Eq. (23).

For multibody final states, Eqs. (13)–(23) apply separately to each point in phase-space. Although  $x$  and  $y$  do not vary across the space, knowledge of the resonant substructure is needed to extrapolate the strong phase difference  $\delta$  from point to point to determine  $x$  and  $y$ . Model-independent methods to measure  $D$  mixing parameters require input related to the relative phases of the  $D^0$  and  $\bar{D}^0$  decay amplitudes across the

# Meson Particle Listings

## $D^0$

phase-space distribution [25]. The required phase information is accessible at the charm threshold, where CLEO-c and BESIII operate [26,27].

A time-dependent analysis of the process  $D^0 \rightarrow K^+\pi^-\pi^0$  from BaBar [28,29] determines the *relative* strong phase variation across the Dalitz plot and reports  $x'' = (2.61^{+0.57}_{-0.68} \pm 0.39)\%$ , and  $y'' = (-0.06^{+0.55}_{-0.64} \pm 0.34)\%$ , where  $x''$  and  $y''$  are defined as

$$\begin{aligned} x'' &\equiv x \cos \delta_{K\pi\pi^0} + y \sin \delta_{K\pi\pi^0}, \\ y'' &\equiv y \cos \delta_{K\pi\pi^0} - x \sin \delta_{K\pi\pi^0}, \end{aligned} \quad (28)$$

in parallel to  $x'$ ,  $y'$ , and  $\delta_{K\pi}$  of Eq. (21). Here  $\delta_{K\pi\pi^0}$  is the remaining strong phase difference between the DCS  $D^0 \rightarrow K^+\rho^-$  and the CF  $\overline{D}^0 \rightarrow K^+\rho^-$  amplitudes and does not vary across the Dalitz plot. Both strong phases,  $\delta_{K\pi}$  and  $\delta_{K\pi\pi^0}$ , can be determined from time-integrated  $CP$  asymmetries in correlated  $D^0\overline{D}^0$  produced at the  $\psi(3770)$  [26,27].

Both the sign and magnitude of  $x$  and  $y$  without phase or sign ambiguity may be measured using the time-dependent resonant substructure of multibody  $D^0$  decays [30,31]. In  $D^0 \rightarrow K_S^0\pi^+\pi^-$ , the DCS and CF decay amplitudes populate the same Dalitz plot, which allows direct measurement of the relative strong phases. CLEO [32], Belle [31,34], and BaBar [33] have measured the relative phase between  $D^0 \rightarrow K^*(892)^-\pi^+$  and  $D^0 \rightarrow K^*(892)^+\pi^-$  to be  $(189 \pm 10 \pm 3^{+15}_{-5})^\circ$ ,  $(173.9 \pm 0.7 \text{ (stat. only)})^\circ$ , and  $(177.6 \pm 1.1 \text{ (stat. only)})^\circ$ , respectively. These results are close to the  $180^\circ$  expected from Cabibbo factors and a small strong phase. Table 4 summarizes the results of a time-dependent Dalitz-plot analyses.

**Table 4:** Results from time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0\pi^+\pi^-$  (CLEO and Belle) and  $D^0 \rightarrow K_S^0\pi^+\pi^-, K_S^0K^+K^-$  (BaBar). The errors are statistical, experimental systematic, and decay-model systematic, respectively.

No $CP$ Violation				
Year	Exper.	$x \times 10^{-3}$	$y \times 10^{-3}$	
2014	Belle [34]	$5.6 \pm 1.9^{+0.3+0.6}_{-0.9-0.9}$	$3.0 \pm 1.5^{+0.4+0.3}_{-0.5-0.8}$	
2010	BaBar [33]	$1.6 \pm 2.3 \pm 1.2 \pm 0.8$	$5.7 \pm 2.0 \pm 1.3 \pm 0.7$	
2007	Belle [31]	$8.0 \pm 2.9^{+0.9+1.0}_{-0.7-1.4}$	$3.3 \pm 2.4^{+0.8+0.6}_{-1.2-0.8}$	
2005	CLEO [30]	$19^{+32}_{-33} \pm 4 \pm 4$	$-14 \pm 24 \pm 8 \pm 4$	
With $CP$ Violation				
Year	Exper.	$ q/p $	$\phi$	
2014	Belle [34]	$0.90^{+0.16+0.05+0.06}_{-0.15-0.04-0.05}$	$(-6 \pm 11 \pm 3^{+3}_{-4})^\circ$	
2007	Belle [31]	$0.86^{+0.30+0.06}_{-0.29-0.03} \pm 0.08$	$(-14^{+16+5+2}_{-18-3-4})^\circ$	

In addition, Belle [31,34] has results for both the relative phase (statistical errors only) and ratio  $R$  (central values only) of the DCS fit fraction relative to the CF fit fractions for  $K^*(892)^+\pi^-$ ,  $K_0^*(1430)^+\pi^-$ ,  $K_2^*(1430)^+\pi^-$ ,  $K^*(1410)^+\pi^-$ , and  $K^*(1680)^+\pi^-$ . Similarly, BaBar [33,35,36] has reported central values for  $R$  for  $K^*(892)^+\pi^-$ ,  $K_0^*(1430)^+\pi^-$ , and

$K_2^*(1430)^+\pi^-$ . The systematic uncertainties on  $R$  must be evaluated. The large differences in  $R$  among these final states could point to an interesting role for hadronic effects.

**Decays to  $CP$  Eigenstates:** When the final state  $f$  is a  $CP$  eigenstate, there is no distinction between  $f$  and  $\overline{f}$ , and  $A_f = A_{\overline{f}}$  and  $\overline{A}_{\overline{f}} = \overline{A}_f$ . We denote final states with  $CP$  eigenvalues  $\pm 1$  by  $f_\pm$  and write  $\lambda_\pm$  for  $\lambda_{f_\pm}$ .

The quantity  $y$  may be measured by comparing the rate for  $D^0$  decays to non- $CP$  eigenstates such as  $K^-\pi^+$  with decays to  $CP$  eigenstates such as  $K^+K^-$  [12]. If decays to  $K^+K^-$  have a shorter effective lifetime than those to  $K^-\pi^+$ ,  $y$  is positive.

In the limit of slow mixing ( $x, y \ll 1$ ) and the absence of direct  $CP$  violation ( $A_D = 0$ ), but allowing for small indirect  $CP$  violation ( $|A_M|, |\phi| \ll 1$ ), we can write

$$\lambda_\pm = \left| \frac{q}{p} \right| e^{\pm i\phi}. \quad (29)$$

In this scenario, to a good approximation, the decay rates for states that are initially  $D^0$  and  $\overline{D}^0$  to a  $CP$  eigenstate have exponential time dependence:

$$r_\pm(t) \propto \exp(-t/\tau_\pm), \quad (30)$$

$$\overline{r}_\pm(t) \propto \exp(-t/\overline{\tau}_\pm), \quad (31)$$

where  $\tau$  is measured in units of  $1/\Gamma$ .

The effective lifetimes are given by

$$1/\tau_\pm = 1 \pm \left| \frac{q}{p} \right| (y \cos \phi - x \sin \phi), \quad (32)$$

$$1/\overline{\tau}_\pm = 1 \pm \left| \frac{p}{q} \right| (y \cos \phi + x \sin \phi). \quad (33)$$

The effective decay rate to a  $CP$  eigenstate combining both  $D^0$  and  $\overline{D}^0$  decays is

$$r_\pm(t) + \overline{r}_\pm(t) \propto e^{-(1 \pm y_{CP})t}. \quad (34)$$

Here

$$y_{CP} = \frac{1}{2} \left( \left| \frac{q}{p} \right| + \left| \frac{p}{q} \right| \right) y \cos \phi - \frac{1}{2} \left( \left| \frac{q}{p} \right| - \left| \frac{p}{q} \right| \right) x \sin \phi \quad (35)$$

$$\approx y \cos \phi - A_M x \sin \phi. \quad (36)$$

If  $CP$  is conserved,  $y_{CP} = y$ .

All measurements of  $y_{CP}$  are relative to the  $D^0 \rightarrow K^-\pi^+$  decay rate. Table 5 summarizes the current status of measurements. Belle [41], BaBar [42], LHCb [43], CDF [39] have reported  $y_{CP}$  and the decay-rate asymmetry for  $CP$  even final states (assuming  $A_D = 0$ )

$$A_\Gamma = \frac{\overline{\tau}_+ - \tau_+}{\overline{\tau}_+ + \tau_+} = \frac{(1/\tau_+) - (1/\overline{\tau}_+)}{(1/\tau_+) + (1/\overline{\tau}_+)} \quad (37)$$

$$= \frac{1}{2} \left( \left| \frac{q}{p} \right| - \left| \frac{p}{q} \right| \right) y \cos \phi - \frac{1}{2} \left( \left| \frac{q}{p} \right| + \left| \frac{p}{q} \right| \right) x \sin \phi \quad (38)$$

$$\approx A_M y \cos \phi - x \sin \phi. \quad (39)$$



**Table 5:** Results for  $y_{CP}$  from  $D^0 \rightarrow K^+K^-$  and  $\pi^+\pi^-$ .

Year	Exper.	final state(s)	$y_{CP}(\%)$	$A_\Gamma(\times 10^{-3})$
2015	LHCb [37]	$K^+K^-, \pi^+\pi^-$	—	$-1.25 \pm 0.73$
2015	LHCb [37]	$K^+K^-$	—	$-1.34 \pm 0.77^{+0.26}_{-0.34}$
2015	LHCb [37]	$\pi^+\pi^-$	—	$-0.092 \pm 1.45^{+0.25}_{-0.33}$
2015	BES III [38]	$K_S^0\pi^0, K_S^0\eta, K_S^0\omega$ $K^+K^-, \pi^+\pi^-$ , $K_S^0\pi^0\pi^0$	$-2.0 \pm 1.3 \pm 0.7$	—
2014	CDF [39]	$K^+K^-, \pi^+\pi^-$	—	$-1.12 \pm 1.2$
2014	CDF [39]	$K^+K^-$	—	$-1.9 \pm 1.5 \pm 0.4$
2014	CDF [39]	$\pi^+\pi^-$	—	$-0.1 \pm 1.8 \pm 0.3$
2013	LHCb [40]	$K^+K^-$	—	$-0.35 \pm 0.62 \pm 0.12$
2013	LHCb [40]	$\pi^+\pi^-$	—	$0.33 \pm 1.06 \pm 0.14$
2012	Belle [41]	$K^+K^-, \pi^+\pi^-$	$1.11 \pm 0.22 \pm 0.11$	$-0.3 \pm 2.0 \pm 0.8$
2012	BaBar [42]	$K^+K^-, \pi^+\pi^-$	$0.72 \pm 0.18 \pm 0.12$	$0.9 \pm 2.6 \pm 0.6$
2011	LHCb [43]	$K^+K^-$	$0.55 \pm 0.63 \pm 0.41$	$-5.9 \pm 5.9 \pm 2.1$
2009*	BaBar [44]	$K^+K^-$	$1.16 \pm 0.22 \pm 0.18$	—
2009	Belle [45]	$K_S^0K^+K^-$	$0.11 \pm 0.61 \pm 0.52$	—
2008*	BaBar [46]	$K^+K^-, \pi^+\pi^-$	$1.03 \pm 0.33 \pm 0.19$	$2.6 \pm 3.6 \pm 0.8$
2007*	Belle [47]	$K^+K^-, \pi^+\pi^-$	$1.31 \pm 0.32 \pm 0.25$	$0.1 \pm 3.0 \pm 1.5$
2003*	BaBar [48]	$K^+K^-, \pi^+\pi^-$	$0.8 \pm 0.4^{+0.5}_{-0.4}$	—
2001	CLEO [49]	$K^+K^-, \pi^+\pi^-$	$-1.2 \pm 2.5 \pm 1.4$	—
2001	Belle† [50]	$K^+K^-$	$-0.5 \pm 1.0^{+0.7}_{-0.8}$	—
2000	FOCUS [51]	$K^+K^-$	$3.42 \pm 1.39 \pm 0.74$	—
1999	E791 [52]	$K^+K^-$	$0.8 \pm 2.9 \pm 1.0$	—
HFAG [24]			$0.835 \pm 0.155$	$-0.59 \pm 0.40$

\*These measurements are excluded from the HFAG average. The BaBar result is superseded by Ref. 42 and the Belle result has been superseded by Ref. 41.

**Table 6:** Results for the difference in time-integrated  $CP$  asymmetry  $\Delta A_{CP}$  between  $D^0 \rightarrow K^+K^-$  and  $D^0 \rightarrow \pi^+\pi^-$ .

Year	Exper.	$\Delta A_{CP}(\times 10^{-3})$
2014	LHCb [54]	$1.4 \pm 1.6 \pm 0.8$
2013	LHCb [55]	$-3.4 \pm 1.5 \pm 1.0$
2013	CDF [56]	$-6.2 \pm 2.1 \pm 1.0$
2012	Belle [14]	$-8.7 \pm 4.1 \pm 0.6$
2008	BaBar [57]	$2.4 \pm 6.2 \pm 2.6$
HFAG [24]		$-2.53 \pm 1.04$

Belle [45] has also reported  $y_{CP}$  for the final state  $K_S^0K^+K^-$  which is dominated by the  $CP$  odd final state  $K_S^0\phi$ . If  $CP$  is conserved,  $A_\Gamma = 0$ .

Substantial work on the time-integrated  $CP$  asymmetries in decays to  $CP$  eigenstates are summarized in this *Review* [53]. Table 6 summarizes the current status of measurements of the difference in time-integrated  $CP$  asymmetry,  $\Delta A_{CP} = A_K - A_\pi$ , between  $D^0 \rightarrow K^+K^-$  and  $D^0 \rightarrow \pi^+\pi^-$ . The HFAG fit is marginally consistent with no  $CP$  violation at the 5.1% Confidence Level [24].

**Coherent  $D^0\bar{D}^0$  Analyses:** Measurements of  $R_D$ ,  $\cos\delta_{K\pi}$ ,  $\sin\delta_{K\pi}$ ,  $x$ , and  $y$  can be determined simultaneously from a combined fit to the time-integrated single-tag (ST) and double-tag (DT) yields in correlated  $D^0\bar{D}^0$  produced at the  $\psi(3770)$  [26,27].

Due to quantum correlations in the  $C = -1$  and  $C = +1$   $D^0\bar{D}^0$  pairs produced in the reactions  $e^+e^- \rightarrow D^0\bar{D}^0(\pi^0)$  and  $e^+e^- \rightarrow D^0\bar{D}^0\gamma(\pi^0)$ , respectively, the time-integrated  $D^0\bar{D}^0$

decay rates are sensitive to interference between amplitudes for indistinguishable final states. The size of this interference is governed by the relevant amplitude ratios and can include contributions from  $D^0\bar{D}^0$  mixing.

The following categories of final states are considered:

**$f$  or  $\bar{f}$ :** Hadronic states accessed from either  $D^0$  or  $\bar{D}^0$  decay but that are not  $CP$  eigenstates. An example is  $K^-\pi^+$ , which results from Cabibbo-favored  $D^0$  transitions or DCS  $\bar{D}^0$  transitions.

**$\ell^+$  or  $\ell^-$ :** Semileptonic or purely leptonic final states, which, in the absence of mixing, tag unambiguously the flavor of the parent  $D^0$ .

**$f_+$  or  $f_-$ :**  $CP$ -even and  $CP$ -odd eigenstates, respectively.

The decay rates for  $D^0\bar{D}^0$  pairs to all possible combinations of the above categories of final states are calculated in Ref. 2, for both  $C = -1$  and  $C = +1$ , reproducing the work of Ref. 3. Such  $D^0\bar{D}^0$  combinations, where both  $D$  final states are specified, are double tags. In addition, the rates for single tags, where either the  $D^0$  or  $\bar{D}^0$  is identified and the other neutral  $D$  decays generically are given in Ref. 2.

BESIII has reported results using  $2.92 \text{ pb}^{-1}$  of  $e^+e^- \rightarrow \psi(3770)$  data where the quantum-coherent  $D^0\bar{D}^0$  pairs are in the  $C = -1$  state. The values of  $y_{CP} = (-2.0 \pm 1.3 \pm 0.7)\%$  [38] and  $A_{K\pi}^{CP} = (12.7 \pm 1.3 \pm 0.7)\%$  [61] are determined from DT yields including a  $CP$  eigenstate vs semileptonic and vs  $K\pi$ , respectively. For  $y_{CP}$ , the  $CP$  eigenstates included are  $K^-K^+$  ( $f_+$ ),  $\pi^+\pi^-$  ( $f_+$ ),  $K_S^0\pi^0\pi^0$  ( $f_+$ ),  $K_S^0\pi^0$  ( $f_-$ ),  $K_S^0\eta$  ( $f_-$ ), and  $K_S^0\omega$  ( $f_-$ ). For  $A_{K\pi}^{CP}$ , the additional  $CP$  eigenstates included are  $\pi^0\pi^0$  ( $f_+$ ) and  $\rho^0\pi^0$  ( $f_+$ ). Using the external inputs from  $R_D$  and  $y$  from HFAG [62] and  $R$  from PDG [63]—see Eq. (27), they obtain  $\cos\delta_{K\pi} = 1.02 \pm 0.11 \pm 0.06 \pm 0.01$  [61] where the third uncertainty is due to the external inputs.

CLEO-c has reported results using  $818 \text{ pb}^{-1}$  of  $e^+e^- \rightarrow \psi(3770)$  data [58–60]. The values of  $y$ ,  $R_M$ ,  $\cos\delta_{K\pi}$ , and  $\sin\delta_{K\pi}$  are determined from a combined fit to the ST (hadronic only) and DT yields. The hadronic final states included are  $K^-\pi^+$  ( $f$ ),  $K^+\pi^-$  ( $\bar{f}$ ),  $K^-K^+$  ( $f_+$ ),  $\pi^+\pi^-$  ( $f_+$ ),  $K_S^0\pi^0\pi^0$  ( $f_+$ ),  $K_L^0\pi^0$  ( $f_+$ ),  $K_L^0\eta$  ( $f_+$ ),  $K_L^0\omega$  ( $f_+$ ),  $K_S^0\pi^0$  ( $f_-$ ),  $K_S^0\eta$  ( $f_-$ ),  $K_S^0\omega$  ( $f_-$ ), and  $K_L^0\pi^0\pi^0$  ( $f_-$ ), and  $K_S^0\pi^+\pi^-$  (mixture of  $f, \bar{f}, f_+$ , and  $f_-$ ). The two flavored final states,  $K^-\pi^+$  and  $K^+\pi^-$ , can be reached via CF or DCS transitions.

Semileptonic DT yields are also included, where one  $D$  is fully reconstructed in one of the hadronic modes listed above, and the other  $D$  is partially reconstructed in either  $D \rightarrow K\ell\nu$  or  $D \rightarrow K\mu\nu$ . When the lepton is accompanied by a flavor tag ( $D \rightarrow K^-\pi^+$  or  $K^+\pi^-$ ), both the “right-sign” and “wrong-sign” DT samples are used, where the electron and kaon charges are the same and opposite, respectively.

The main results of the CLEO-c analysis are the determination of  $\cos\delta_{K\pi} = 0.81^{+0.22+0.07}_{-0.18-0.05}$ ,  $\sin\delta_{K\pi} = -0.01 \pm 0.49 \pm 0.04$ , and World Averages for the mixing parameters from an “extended” fit that combines the CLEO-c data with previous mixing and branching-ratio measurements [60]. These fits allow  $\cos\delta_{K\pi}$ ,  $\sin\delta_{K\pi}$  and  $x^2$  to be unphysical. Constraining

# Meson Particle Listings

## $D^0$

$\cos \delta_{K\pi}$  and  $\sin \delta_{K\pi}$  to  $[-1, +1]$ —that is interpreting  $\delta_{K\pi}$  as an angle—yields  $\delta_{K\pi} = (18^{+11}_{-17} \pm 7)^\circ$ . Note that measurements of  $y$  (Table 4 and Table 5) and  $y'$  (Table 3) contribute to the determination of  $\delta_{K\pi}$ .

**Summary of Experimental Results:** Several recent results indicate that charm mixing is at the upper end of the range of Standard Model estimates.

For  $D^0 \rightarrow K^+\pi^-$ , LHCb [15,17], CDF [16], and Belle [14] each exclude the no-mixing hypothesis by more than 5 standard deviations.

For  $y_{CP}$  in  $D^0 \rightarrow K^+K^-$  and  $\pi^+\pi^-$ , Belle [41] and BaBar [42] find  $4.5\sigma$  and  $3.3\sigma$  effects. The most sensitive measurement of  $x$  and  $y$  is in  $D^0 \rightarrow K_S^0\pi^+\pi^-$  from Belle [34] and the no mixing solution is only excluded at  $2.5\sigma$ . In a similar analysis using  $D^0 \rightarrow K_S^0\pi^+\pi^-$  and  $D^0 \rightarrow K_S^0K^+K^-$  BaBar [33] also finds the no mixing solution excluded at  $1.9\sigma$ .

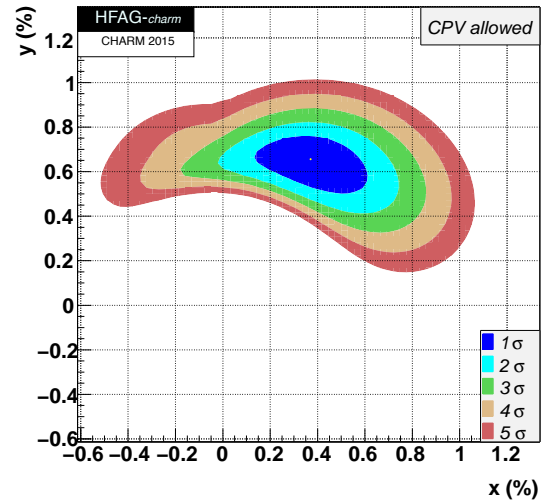
The current situation would benefit from better knowledge of the strong phase difference  $\delta_{K\pi}$  than provided by the current CLEO-c [60] and BESIII [61] results. This would allow one to unfold  $x$  and  $y$  from the  $D^0 \rightarrow K^+\pi^-$  measurements of  $x'^2$  and  $y'$ , and directly compare them to the  $D^0 \rightarrow K_S^0\pi^+\pi^-$  results.

The experimental data consistently indicate that the  $D^0$  and  $\bar{D}^0$  do mix. The mixing is presumably dominated by long-range processes. Under the assumption that the observed mixing is due entirely to non-Standard Model processes, significant constraints on a variety of new physics models are obtained [64]. A serious limitation to the interpretation of charm oscillations in terms of New Physics is the theoretical uncertainty of the Standard Model prediction. The evidence for time integrated  $CP$ -violation,  $\Delta A_{CP} \neq 0$  is intriguing. This result is marginally consistent with Standard Model expectation [65–67].

### HFAG Averaging of Charm Mixing Results:

The Heavy Flavor Averaging Group (HFAG) has made a global fit to all mixing measurements to obtain values of  $x$ ,  $y$ ,  $\delta_{K\pi}$ ,  $\delta_{K\pi\pi^0}$ ,  $R_D$ ,  $A_D \equiv (R_D^+ - R_D^-)/(R_D^+ + R_D^-)$ ,  $|q/p|$ ,  $\text{Arg}(q/p) \equiv \phi$ , and the time-integrated  $CP$  asymmetries  $A_K$  and  $A_\pi$ . Correlations among observables are taken into account by using the error matrices from the experiments. The measurements of  $D^0 \rightarrow K^{(*)}\ell^-\bar{\nu}$ ,  $K^+K^-$ ,  $\pi^+\pi^-$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+\pi^-$ ,  $K_S^0\pi^+\pi^-$ , and  $K_S^0K^+K^-$  decays, as well as CLEO-c and BESIII results for double-tagged branching fractions measured at the  $\psi(3770)$  are used.

For the global fit, confidence contours in the two dimensions  $(x, y)$  and  $(|q/p|, \phi)$  are obtained by letting, for any point in the two-dimensional plane, all other fit parameters take their preferred values. Figures 1 and 2 show the resulting 1-to-5  $\sigma$  contours. The fits exclude the no-mixing point ( $x = y = 0$ ) at more than  $11.5\sigma$ , when  $CP$  violation is allowed. The fits are consistent with no  $CP$  violation at the 27% Confidence Level. The parameters  $x$  and  $y$  differ from zero by  $2.1\sigma$  and  $6.8\sigma$ , respectively. One-dimensional likelihood functions for parameters are obtained by allowing, for any value of the parameter, all other fit parameters to take their preferred values. The



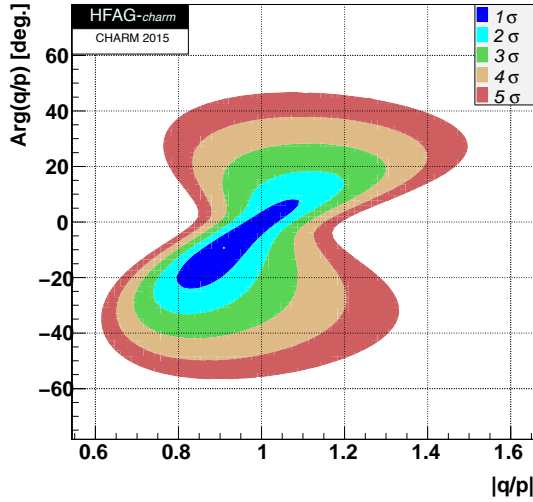
**Figure 1:** Two-dimensional  $1\sigma$ - $5\sigma$  contours for  $(x, y)$  from measurements of  $D^0 \rightarrow K^{(*)}\ell\nu$ ,  $h^+h^-$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+\pi^-$ ,  $K_S^0\pi^+\pi^-$ , and  $K_S^0K^+K^-$  decays, and double-tagged branching fractions measured at the  $\psi(3770)$  resonance (from HFAG [24]).

**Table 7:** HFAG Charm Mixing Averages [24].

Parameter	No $CP$ Violation	$CP$ Violation Allowed	95% C.L. Interval
$x(\%)$	$0.49^{+0.14}_{-0.15}$	$0.37 \pm 0.16$	$[0.06, 0.67]$
$y(\%)$	$0.61 \pm 0.08$	$0.66^{+0.07}_{-0.10}$	$[0.46, 0.79]$
$R_D(\%)$	$0.349 \pm 0.004$	$0.349 \pm 0.004$	$[0.342, 0.357]$
$\delta_{K\pi}(\circ)$	$6.9^{+9.7}_{-11.2}$	$11.8^{+9.5}_{-14.7}$	$[-21.1, 29.3]$
$\delta_{K\pi\pi^0}(\circ)$	$18.1^{+23.2}_{-23.8}$	$27.3^{+24.4}_{-25.4}$	$[-23.3, 74.8]$
$A_D(\%)$	—	$-0.39^{+1.01}_{-1.05}$	$[-2.4, 1.5]$
$ q/p $	—	$0.91^{+0.12}_{-0.08}$	$[0.77, 1.14]$
$\phi(\circ)$	—	$-9.4^{+11.9}_{-9.8}$	$[-28.3, 12.9]$
$A_K$	—	$-0.15 \pm 0.14$	$[-0.42, 0.12]$
$A_\pi$	—	$0.10 \pm 0.15$	$[-0.19, 0.38]$

resulting likelihood functions give central values, 68.3% C.L. intervals, and 95% C.L. intervals as listed in Table 7. The  $\chi^2$  for the HFAG fit is 69 for 45 degrees of freedom indicating some disagreement among the measurements included in the combination.

From the results of the HFAG averaging, the following can be concluded: (1) Since  $CP$  violation is small and  $y_{CP}$  is positive, the  $CP$ -even state is shorter-lived, as in the  $K^0\bar{K}^0$  system; (2) However, since  $x$  appears to be positive, the  $CP$ -even state is heavier, unlike in the  $K^0\bar{K}^0$  system; (3) The strong phase difference  $\delta_{K\pi}$  is consistent with the SU(3) expectation of zero but large values are not excluded; (4) There is no evidence yet for  $CP$ -violation in  $D^0\bar{D}^0$  mixing. Observing  $CP$ -violation in mixing ( $|q/p| \neq 1$ ) at the current level of sensitivity would indicate new physics.



**Figure 2:** Two-dimensional  $1\sigma$ - $5\sigma$  contours for  $(|q/p|, \text{Arg}(q/p))$  from measurements of  $D^0 \rightarrow K^{(*)+} \ell \nu, h^+ h^-, K^+ \pi^-, K^+ \pi^- \pi^0, K^+ \pi^- \pi^+ \pi^-, K_S^0 \pi^+ \pi^-,$  and  $K_S^0 K^+ K^-$  decays, and double-tagged branching fractions measured at the  $\psi(3770)$  resonance (from HFAG [24]).

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## References

1. B. Aubert *et al.*, Phys. Rev. **D76**, 014018 (2007).
2. D.M. Asner and W.M. Sun, Phys. Rev. **D73**, 034024 (2006); Erratum-*ibid.*, **77**, 019901 (2008).
3. D. Atwood and A.A. Petrov, Phys. Rev. **D71**, 054032 (2005), M. Gronau, Y. Grossman, J.L. Rosner Phys. Lett. **B508**, 37 (2001), Z.Z. Xing, Phys. Rev. **D55**, 196 (1997), M. Goldhaber and J.L. Rosner, Phys. Rev. **D15**, 1254 (1977).
4. U. Bitenc *et al.*, Phys. Rev. **D77**, 112003 (2008).
5. U. Bitenc *et al.*, Phys. Rev. **D72**, 071101R (2005).
6. C. Cawfield *et al.*, Phys. Rev. **D71**, 077101 (2005).
7. B. Aubert *et al.*, Phys. Rev. **D70**, 091102R (2004).
8. K. Stenson, presented at the April Meeting of the American Physical Society (APS 03), Philadelphia, Pennsylvania, April 5-8, 2003; M. Hosack, (FOCUS Collab.), Fermilab-Thesis-2002-25.
9. E.M. Aitala *et al.*, Phys. Rev. Lett. **77**, 2384 (1996).
10. A.J. Schwartz, for the Heavy Flavor Averaging Group, arXiv:0911.1464 [hep-ex].
11. Y. Nir, Lectures given at 27th SLAC Summer Institute on Particle Physics: “ $CP$  Violation in and Beyond the Standard Model (SSI 99),” Stanford, California, 7-16 Jul 1999. Published in Trieste 1999, *Particle Physics*, pp. 165-243.
12. S. Bergmann *et al.*, Phys. Lett. **B486**, 418 (2000).
13. See the Note on “ $CP$  Violation in Meson Decays” in this Review.
14. B.R. Ko *et al.*, Phys. Rev. Lett. **112**, 111801 (2014).
15. R. Aaij *et al.*, Phys. Rev. Lett. **111**, 251801 (2013).
16. T. Aaltonen *et al.*, Phys. Rev. Lett. **111**, 231802 (2013).
17. R. Aaij *et al.*, Phys. Rev. Lett. **110**, 101802 (2013).
18. T. Aaltonen *et al.*, Phys. Rev. Lett. **100**, 121802 (2008).
19. B. Aubert *et al.*, Phys. Rev. Lett. **98**, 211802 (2007).
20. L.M. Zhang *et al.*, Phys. Rev. Lett. **96**, 151801 (2006).
21. J.M. Link *et al.*, Phys. Lett. **B607**, 51 (2005).
22. R. Godang *et al.*, Phys. Rev. Lett. **84**, 5038 (2000).
23. E.M. Aitala *et al.*, Phys. Rev. **D57**, 13 (1998).
24. Heavy Flavor Averaging Group, [slac.stanford.edu/xorg/hfag/charm/CHARM15/results\\_mix\\_cpv.html](http://slac.stanford.edu/xorg/hfag/charm/CHARM15/results_mix_cpv.html).
25. See “Review of Multibody Charm Analyses” in this Review.
26. R.A. Briere *et al.*, (CLEO Collab.), CLNS 01-1742, (2001).
27. G. Cavoto *et al.*, Prepared for *3rd Workshop on the Unitarity Triangle: CKM 2005*, San Diego, California, 15-18 Mar 2005, hep-ph/0603019.
28. B. Aubert *et al.*, Phys. Rev. Lett. **97**, 221803 (2006).
29. B. Aubert *et al.*, Phys. Rev. Lett. **103**, 211801 (2009).
30. D.M. Asner *et al.*, Phys. Rev. **D72**, 012001 (2005).
31. L.M. Zhang *et al.*, Phys. Rev. Lett. **99**, 131803 (2007).
32. H. Muramatsu *et al.*, Phys. Rev. Lett. **89**, 251802 (2002).
33. P. del Amo Sanchez *et al.*, Phys. Rev. Lett. **105**, 081803 (2010).
34. T. Peng *et al.*, Phys. Rev. **D89**, 091103R (2014).
35. B. Aubert *et al.*, Phys. Rev. Lett. **95**, 121802 (2005).
36. B. Aubert *et al.*, Phys. Rev. **D78**, 034023 (2008).
37. R. Aaij *et al.*, JHEP **1504**, 043 (2015).
38. M. Ablikim *et al.*, Phys. Lett. **B744**, 339 (2015).
39. T. Aaltonen *et al.*, Phys. Rev. **D90**, 111103R (2014).
40. R. Aaij *et al.*, Phys. Rev. Lett. **112**, 041801 (2014).
41. M. Staric *et al.*, arXiv:1212.3478[hep-ex].
42. B. Aubert *et al.*, Phys. Rev. **D87**, 012004 (2013).
43. R. Aaij *et al.*, JHEP **1204**, 129 (2012).
44. B. Aubert *et al.*, Phys. Rev. **D80**, 071103R (2009).
45. A. Zupanc *et al.*, Phys. Rev. **D80**, 052006 (2009).
46. B. Aubert *et al.*, Phys. Rev. **D78**, 011105 (2008).
47. M. Staric *et al.*, Phys. Rev. Lett. **98**, 211803 (2007).
48. B. Aubert *et al.*, Phys. Rev. Lett. **91**, 121801 (2003).
49. S.E. Csorna *et al.*, Phys. Rev. **D65**, 092001 (2002).
50. K. Abe *et al.*, Phys. Rev. Lett. **88**, 162001 (2002).
51. J.M. Link *et al.*, Phys. Lett. **B485**, 62 (2000).
52. E.M. Aitala *et al.*, Phys. Rev. Lett. **83**, 32 (1999).
53. See the tabulation of  $A_{CP}$  results in the  $D^0$  and  $D^+$  Listings in this Review.
54. R. Aaij *et al.*, JHEP **1407**, 041 (2014).
55. R. Aaij *et al.*, LHCb-CONF-2013-003.
56. T. Aaltonen *et al.*, Phys. Rev. Lett. **109**, 111801 (2012).
57. B. Aubert *et al.*, Phys. Rev. Lett. **100**, 061803 (2008).
58. J.L. Rosner *et al.*, Phys. Rev. Lett. **100**, 221801 (2008).
59. D.M. Asner *et al.*, Phys. Rev. **D78**, 012001 (2008).
60. D.M. Asner *et al.*, Phys. Rev. **D86**, 112001 (2012).
61. M. Ablikim *et al.*, Phys. Lett. **B734**, 227 (2014).
62. Heavy Flavor Averaging Group, [slac.stanford.edu/xorg/hfag/charm/PDG14/results\\_mix\\_cpv.html](http://slac.stanford.edu/xorg/hfag/charm/PDG14/results_mix_cpv.html).
63. J. Beringer *et al.*, Phys. Rev. **D86**, 010001 (2012).

# Meson Particle Listings

$D^0$

64. E. Golowich *et al.*, Phys. Rev. **D76**, 095009 (2007).
65. G. Isidori *et al.*, Phys. Lett. **B711**, 46 (2011).
66. E. Franco *et al.*, JHEP **1205**, 140 (2012).
67. M. Gersaback *et al.*, J. Phys. **G39**, 045005 (2012).

$$|m_{D_1^0} - m_{D_2^0}| = x \Gamma$$

The  $D_1^0$  and  $D_2^0$  are the mass eigenstates of the  $D^0$  meson, as described in the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. The experiments usually present  $x \equiv \Delta m/\Gamma$ . Then  $\Delta m = x \Gamma = x \hbar/\tau$ .

“OUR EVALUATION” comes from CPV allowing averages provided by the Heavy Flavor Averaging Group, see the note on “ $D^0$ - $\bar{D}^0$  Mixing.”

VALUE ( $10^{10} \hbar s^{-1}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.95<sup>+0.41</sup><sub>-0.44</sub></b>		<b>OUR EVALUATION</b>		
<b>1.0 ± 0.5</b>		<b>OUR AVERAGE</b>		
		Error includes scale factor of 1.2.		
		1 KO	14 BELL	$e^+ e^- \rightarrow \Upsilon(nS)$
1.37 ± 0.46 <sup>+0.18</sup> <sub>-0.28</sub>		2 PENG	14 BELL	$e^+ e^- \rightarrow \Upsilon(nS)$
		3 AAIJ	13CE LHCb	$p\bar{p}$ at 7, 8 TeV
		4 AALTONEN	13AE CDF	$p\bar{p}$ at 1.96 TeV
0.39 ± 0.56 ± 0.35		5 DEL-AMO-SA...10D	BABR	$e^+ e^-$ , 10.6 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		6 AAIJ	13N LHCb	Repl. by AAIJ 13CE
6.4 <sup>+1.4</sup> <sub>-1.7</sub> ± 1.0		7 AUBERT	09AN BABR	$e^+ e^-$ at 10.58 GeV
– 2 <sup>+7</sup> <sub>-6</sub>		8 LOWREY	09 CLEO	$e^+ e^-$ at $\psi(3770)$
1.98 ± 0.73 <sup>+0.32</sup> <sub>-0.41</sub>		9 ZHANG	07B BELL	Repl. by PENG 14
< 7	95	10 ZHANG	06 BELL	$e^+ e^-$
– 11 to +22		9 ASNER	05 CLEO	$e^+ e^- \approx 10$ GeV
< 11	90	BITENC	05 BELL	
< 30	90	CAWLFIELD	05 CLEO	
< 7	95	10 LI	05A BELL	See ZHANG 06
< 22	95	11 LINK	05H FOCS	$\gamma$ nucleus
< 23	95	AUBERT	04Q BABR	
< 11	95	10 AUBERT	03Z BABR	$e^+ e^-$ , 10.6 GeV
< 7	95	12 GODANG	00 CLE2	$e^+ e^-$
< 32	90	13,14 AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
< 24	90	15 AITALA	96C E791	$\pi^-$ nucleus, 500 GeV
< 21	90	14,16 ANJOS	88C E691	Photoproduction

<sup>1</sup> Based on  $976 \text{ fb}^{-1}$  of data collected at  $Y(nS)$  resonances. Assumes no  $CP$  violation. Reported  $x'^2 = (0.09 \pm 0.22) \times 10^{-3}$  and  $y' = (4.6 \pm 3.4) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$ .

<sup>2</sup> The time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$  is employed. Decay-time information and interference on the Dalitz plot are used to distinguish doubly Cabibbo-suppressed decays from mixing and to measure the relative phase between  $D^0 \rightarrow K^{*+} \pi^-$  and  $\bar{D}^0 \rightarrow K^{*+} \pi^-$ . This value allows  $CP$  violation and is sensitive to the sign of  $\Delta m$ .

<sup>3</sup> Based on  $3 \text{ fb}^{-1}$  of data collected at  $\sqrt{s} = 7, 8 \text{ TeV}$ . Assumes no  $CP$  violation. Reported  $x'^2 = (5.5 \pm 4.9) \times 10^{-4}$  and  $y' = (4.8 \pm 1.0) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between the  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$ .

<sup>4</sup> Based on  $9.6 \text{ fb}^{-1}$  of data collected at the Tevatron. Assumes no  $CP$  violation. Reported  $x'^2 = (0.08 \pm 0.18) \times 10^{-3}$  and  $y' = (4.3 \pm 4.3) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between the  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$ .

<sup>5</sup> DEL-AMO-SANCHEZ 10D uses  $540,800 \pm 800 K_S^0 \pi^+ \pi^-$  and  $79,900 \pm 300 K_S^0 K^+ K^-$  events in a time-dependent amplitude analysis of the  $D^0$  and  $\bar{D}^0$  Dalitz plots. No evidence was found for  $CP$  violation, and the values here assume no such violation.

<sup>6</sup> Based on  $1 \text{ fb}^{-1}$  of data collected at  $\sqrt{s} = 7 \text{ TeV}$  in 2011. Assumes no  $CP$  violation. Reported  $x'^2 = (-0.9 \pm 1.3) \times 10^{-4}$  and  $y' = (7.2 \pm 2.4) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between the  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$ .

<sup>7</sup> The AUBERT 09AN values are inferred from the branching ratio  $\Gamma(D^0 \rightarrow K^+ \pi^- \pi^0)/\Gamma(D^0 \rightarrow K^- \pi^+ \pi^0)$  given near the end of this Listings. Mixing is distinguished from DCS decays using decay-time information. Interference between mixing and DCS is allowed. The phase between  $D^0 \rightarrow K^+ \pi^- \pi^0$  and  $\bar{D}^0 \rightarrow K^+ \pi^- \pi^0$  is assumed to be small. The width difference here is  $y''$ , which is not the same as  $y_{CP}$  in the note on  $D^0$ - $\bar{D}^0$  mixing.

<sup>8</sup> LOWREY 09 uses quantum correlations in  $e^+ e^- \rightarrow D^0 \bar{D}^0$  at the  $\psi(3770)$ . See below for coherence factors and average relative strong phases for both  $D^0 \rightarrow K^- \pi^+ \pi^0$  and  $D^0 \rightarrow K^- \pi^- \pi^+$ . A fit that includes external measurements of charm mixing parameters gets  $\Delta m = (2.34 \pm 0.61) \times 10^{10} \hbar s^{-1}$ .

<sup>9</sup> The ASNER 05 and ZHANG 07B values are from the time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$ . Decay-time information and interference on the Dalitz plot are used to distinguish doubly Cabibbo-suppressed decays from mixing and to measure the relative phase between  $D^0 \rightarrow K^{*+} \pi^-$  and  $\bar{D}^0 \rightarrow K^{*+} \pi^-$ . This value allows  $CP$  violation and is sensitive to the sign of  $\Delta m$ .

<sup>10</sup> The AUBERT 03Z, LI 05A, and ZHANG 06 limits are inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-)/\Gamma(K^- \pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0$ - $\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. AUBERT 03Z assumes the strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  amplitudes is small; if an arbitrary phase is allowed, the limit degrades by 20%. The LI 05A and ZHANG 06 limits are valid for an arbitrary strong phase.

<sup>11</sup> This LINK 05H limit is inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-)$  (via  $\bar{D}^0$ )/ $\Gamma(K^- \pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0$ - $\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. The strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  is assumed to be small. If an arbitrary relative strong phase is allowed, the limit degrades by 25%.

<sup>12</sup> This GODANG 00 limit is inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-)$  (via  $\bar{D}^0$ )/ $\Gamma(K^- \pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0$ - $\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. The strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  is assumed to be small. If an arbitrary relative strong phase is allowed, the limit degrades by a factor of two.

<sup>13</sup> AITALA 98 allows interference between the doubly Cabibbo-suppressed and mixing amplitudes, and also allows  $CP$  violation in this term, but assumes that  $A_D = A_R = 0$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<sup>14</sup> This limit is inferred from  $R_M$  for  $f = K^+ \pi^-$  and  $f = K^+ \pi^- \pi^+ \pi^-$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. Decay-time information is used to distinguish doubly Cabibbo-suppressed decays from  $D^0$ - $\bar{D}^0$  mixing.

<sup>15</sup> This limit is inferred from  $R_M$  for  $f = K^+ \ell^- \pi_\ell$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<sup>16</sup> ANJOS 88c assumes that  $y = 0$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. Without this assumption, the limit degrades by about a factor of two.

$$(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma = 2y$$

The  $D_1^0$  and  $D_2^0$  are the mass eigenstates of the  $D^0$  meson, as described in the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

Due to the strong phase difference between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$ , we exclude from the average those measurements of  $y'$  that are inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-)$  (via  $\bar{D}^0$ ) /  $\Gamma(K^+ \pi^-)$  given near the end of this  $D^0$  Listings.

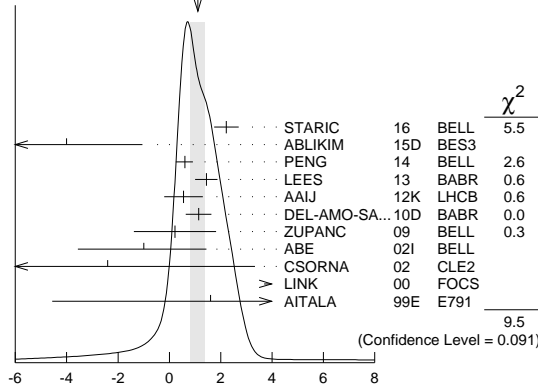
Some early results have been omitted. See our 2006 *Review* (Journal of Physics **G33** 1 (2006)).

“OUR EVALUATION” comes from CPV allowing averages provided by the Heavy Flavor Averaging Group, see the note on “ $D^0$ - $\bar{D}^0$  Mixing.”

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.29<sup>+0.14</sup><sub>-0.18</sub></b>		<b>OUR EVALUATION</b>		
<b>1.11 ± 0.27</b>		<b>OUR AVERAGE</b>		
		Error includes scale factor of 1.4. See the ideogram below.		
2.22 ± 0.44 ± 0.18		1 STARIC	16 BELL	$e^+ e^- \rightarrow \Upsilon(nS)$
–4.0 ± 2.6 ± 1.4		2 ABLIKIM	15D BES3	$e^+ e^-$ at $\psi(3770)$
		3 KO	14 BELL	$e^+ e^- \rightarrow \Upsilon(nS)$
		4 PENG	14 BELL	$e^+ e^- \rightarrow \Upsilon(nS)$
		5 AAIJ	13CE LHCb	$p\bar{p}$ at 7, 8 TeV
		6 AALTONEN	13AE CDF	$p\bar{p}$ at 1.96 TeV
		7 LEES	13 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
		8 AAIJ	12K LHCb	$p\bar{p}$ at 7 TeV
		9 DEL-AMO-SA...10D	BABR	$e^+ e^-$ , 10.6 GeV
0.60 ± 0.30 <sup>+0.10</sup> <sub>-0.17</sub>		10 ZUPANC	09 BELL	$e^+ e^- \approx \Upsilon(4S)$
		11 ABE	02i BELL	$e^+ e^- \approx \Upsilon(4S)$
1.44 ± 0.36 ± 0.24		12 CSORNA	02 CLE2	$e^+ e^- \approx \Upsilon(4S)$
0.55 ± 0.63 ± 0.41		11 LINK	00 FOCS	$\gamma$ nucleus
1.14 ± 0.40 ± 0.30		11 AITALA	99E E791	$K^- \pi^+$ , $K^+ K^-$
0.22 ± 1.22 ± 1.04		• • • We do not use the following data for averages, fits, limits, etc. • • •		
–1.0 ± 2.0 <sup>+1.4</sup> <sub>-1.6</sub>	18k	13 AAIJ	13N LHCb	Repl. by AAIJ 13CE
–2.4 ± 5.0 ± 2.8	3393	14 AUBERT	09Ai BABR	See LEES 13
6.84 ± 2.78 ± 1.48	10k	15 AUBERT	09AN BABR	$e^+ e^-$ at 10.58 GeV
+1.6 ± 5.8 ± 2.1		16 LOWREY	09 CLEO	$e^+ e^-$ at $\psi(3770)$
2.32 ± 0.44 ± 0.36		17 AALTONEN	08E CDF	$p\bar{p}$ , $\sqrt{s} = 1.96 \text{ TeV}$
–0.12 <sup>+1.10</sup> <sub>-1.28</sub> ± 0.68		18 AUBERT	08u BABR	See AUBERT 09Ai
1.4 <sup>+4.8</sup> <sub>-5.4</sub>		17 AUBERT	07W BABR	$e^+ e^- \approx 10.6 \text{ GeV}$
1.70 ± 1.52	12.7 ± 0.3k	19 STARIC	07 BELL	Repl. by STARIC 16
2.06 ± 0.66 ± 0.38		20 ZHANG	07B BELL	Repl. by PENG 14
1.94 ± 0.88 ± 0.62	4030 ± 90	17,21 ZHANG	06 BELL	$e^+ e^-$
2.62 ± 0.64 ± 0.50	160k	20 ASNER	05 CLEO	$e^+ e^- \approx 10 \text{ GeV}$
0.74 ± 0.50 <sup>+0.20</sup> <sub>-0.31</sub>	534k	17,21 LI	05A BELL	See ZHANG 06
–0.7 ± 4.9	4k ± 88	17,21 LINK	05H FOCS	$\gamma$ nucleus
–3.0 ± 5.0 <sup>+1.6</sup> <sub>-4.8</sub> ± 0.8		22 AUBERT	03P BABR	See AUBERT 08u
–0.3 ± 5.7		17,21 AUBERT	03Z BABR	$e^+ e^-$ , 10.6 GeV
–5.2 ± 18.4 <sub>-16.8</sub>		17 GODANG	00 CLE2	$e^+ e^-$
1.6 ± 0.8 <sup>+1.0</sup> <sub>-0.8</sub>	450k			
1.6 ± 6.2 <sub>-12.8</sub>				
–5.0 ± 2.8 ± 0.6				

See key on page 601

## Meson Particle Listings

 $D^0$ WEIGHTED AVERAGE  
1.11±0.27 (Error scaled by 1.4)

$$(\Gamma_1 - \Gamma_2)/\Gamma = 2\gamma$$

- 1 An improved measurement of  $\overline{D}^0 - D^0$  mixing and a search for  $CP$  violation in  $D^0$  decays to  $CP$ -even final states  $K^+ K^-$  and  $\pi^+ \pi^-$  using the final Belle data sample of 976 fb $^{-1}$ .
- 2 ABLIKIM 15D uses quantum correlations in  $e^+ e^- \rightarrow D^0 \overline{D}^0$  at the  $\psi(3770)$ .
- 3 Based on 976 fb $^{-1}$  of data collected at  $Y(nS)$  resonances. Assumes no  $CP$  violation. Reported  $x'^2 = (0.09 \pm 0.22) \times 10^{-3}$  and  $y' = (4.6 \pm 3.4) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\overline{D}^0 \rightarrow K^+ \pi^-$ .
- 4 The time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$  is employed. Decay-time information and interference on the Dalitz plot are used to distinguish doubly Cabibbo-suppressed decays from mixing and to measure the relative phase between  $D^0 \rightarrow K^{*+} \pi^-$  and  $\overline{D}^0 \rightarrow K^{*+} \pi^-$ . This value allows  $CP$  violation and is sensitive to the sign of  $\Delta m$ .
- 5 Based on 3 fb $^{-1}$  of data collected at  $\sqrt{s} = 7, 8$  TeV. Assumes no  $CP$  violation. Reported  $x'^2 = (5.5 \pm 4.9) \times 10^{-4}$  and  $y' = (4.8 \pm 1.0) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between the  $D^0 \rightarrow K^+ \pi^-$  and  $\overline{D}^0 \rightarrow K^+ \pi^-$ .
- 6 Based on 9.6 fb $^{-1}$  of data collected at the Tevatron. Assumes no  $CP$  violation. Reported  $x'^2 = (0.08 \pm 0.18) \times 10^{-3}$  and  $y' = (4.3 \pm 4.3) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between the  $D^0 \rightarrow K^+ \pi^-$  and  $\overline{D}^0 \rightarrow K^+ \pi^-$ .
- 7 Obtained  $y_{CP} = (0.72 \pm 0.18 \pm 0.12)\%$  based on three effective  $D^0$  lifetimes measured in  $K^+ \pi^-$ ,  $K^- K^+$ , and  $\pi^- \pi^+$ . We list  $2y_{CP} = \Delta\Gamma/\Gamma$ .
- 8 Compared the lifetimes of  $D^0$  decay to the  $CP$  eigenstate  $K^+ K^-$  with  $D^0$  decay to  $\pi^+ \pi^-$ . The values here assume no  $CP$  violation.
- 9 DEL-AMO-SANCHEZ 10D uses 540,800  $\pm 800$   $K_S^0 \pi^+ \pi^-$  and 79,900  $\pm 300$   $K_S^0 K^+ K^-$  events in a time-dependent amplitude analyses of the  $D^0$  and  $\overline{D}^0$  Dalitz plots. No evidence was found for  $CP$  violation, and the values here assume no such violation.
- 10 ZUPANC 09 uses a method based on measuring the mean decay time of  $D^0 \rightarrow K_S^0 K^+ K^-$  events for different  $K^+ K^-$  mass intervals.
- 11 LINK 00, AITALA 99E, and ABE 021 measure the lifetime difference between  $D^0 \rightarrow K^- K^+$  ( $CP$ -even) decays and  $D^0 \rightarrow K^- \pi^+$  ( $CP$ -mixed) decays, or  $y_{CP} = [\Gamma(CP^+) - \Gamma(CP^-)] / [\Gamma(CP^+) + \Gamma(CP^-)]$ . We list  $2y_{CP} = \Delta\Gamma/\Gamma$ .
- 12 CSORNA 02 measures the lifetime difference between  $D^0 \rightarrow K^- K^+$  and  $\pi^- \pi^+$  ( $CP$ -even) decays and  $D^0 \rightarrow K^- \pi^+$  ( $CP$ -mixed) decays, or  $y_{CP} = [\Gamma(CP^+) - \Gamma(CP^-)] / [\Gamma(CP^+) + \Gamma(CP^-)]$ . We list  $2y_{CP} = \Delta\Gamma/\Gamma$ .
- 13 Based on 1 fb $^{-1}$  of data collected at  $\sqrt{s} = 7$  TeV in 2011. Assumes no  $CP$  violation. Reported  $x'^2 = (-0.9 \pm 1.3) \times 10^{-4}$  and  $y' = (7.2 \pm 2.4) \times 10^{-3}$ , where  $x' = x \cos(\delta) + y \sin(\delta)$ ,  $y' = y \cos(\delta) - x \sin(\delta)$  and  $\delta$  is the strong phase between the  $D^0 \rightarrow K^+ \pi^-$  and  $\overline{D}^0 \rightarrow K^+ \pi^-$ .
- 14 This combines the  $y_{CP} = (\tau_{K^+ \pi^+} / \tau_{K^+ K^+}) - 1$  using untagged  $K^- \pi^+$  and  $K^- K^+$  events of AUBERT 09AI with the disjoint  $y_{CP}$  using tagged  $K^- \pi^+$ ,  $K^- K^+$ , and  $\pi^- \pi^+$  events of AUBERT 08U.
- 15 The AUBERT 09AN values are inferred from the branching ratio  $\Gamma(D^0 \rightarrow K^+ \pi^- \pi^0)$  via  $\overline{D}^0 / \Gamma(D^0 \rightarrow K^- \pi^+ \pi^0)$  given near the end of this Listings. Mixing is distinguished from DCS decays using decay-time information. Interference between mixing and DCS is allowed. The phase between  $D^0 \rightarrow K^+ \pi^- \pi^0$  and  $\overline{D}^0 \rightarrow K^+ \pi^- \pi^0$  is assumed to be small. The width difference here is  $y'$ , which is not the same as  $y_{CP}$  in the note on  $D^0 - \overline{D}^0$  mixing.
- 16 LOWREY 09 uses quantum correlations in  $e^+ e^- \rightarrow D^0 \overline{D}^0$  at the  $\psi(3770)$ . See below for coherence factors and average relative strong phases for both  $D^0 \rightarrow K^- \pi^+ \pi^0$  and  $D^0 \rightarrow K^- \pi^- 2\pi^+$ . A fit that includes external measurements of charm mixing parameters gets  $2y = (1.62 \pm 0.32) \times 10^{-2}$ .
- 17 The GODANG 00, AUBERT 03Z, LINK 05H, LI 05A, ZHANG 06, AUBERT 07W, and AALTONEN 08E limits are inferred from the  $D^0 - \overline{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-)$  (via  $\overline{D}^0) / \Gamma(K^- \pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0 - \overline{D}^0$  mixing. The limits allow interference between the DCS and mixing ratios, and all except AUBERT 07W and AALTONEN 08E also allow  $CP$  violation. The phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\overline{D}^0 \rightarrow K^+ \pi^-$  is assumed to be small. This is a measurement of  $y'$  and is not the same as the  $y_{CP}$  of our note above on " $D^0 - \overline{D}^0$  Mixing."
- 18 This value combines the results of AUBERT 08U and AUBERT 03P.

- 19 STARIC 07 compares the lifetimes of  $D^0$  decay to the  $CP$  eigenstates  $K^+ K^-$  and  $\pi^+ \pi^-$  with  $D^0$  decay to  $K^- \pi^+$ .
- 20 The ASNER 05 and ZHANG 07B values are from the time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$ . Decay-time information and interference on the Dalitz plot are used to distinguish doubly Cabibbo-suppressed decays from mixing and to measure the relative phase between  $D^0 \rightarrow K^{*+} \pi^-$  and  $\overline{D}^0 \rightarrow K^{*+} \pi^-$ . This limit allows  $CP$  violation.
- 21 The ranges of AUBERT 03Z, LINK 05H, LI 05A, and ZHANG 06 measurements are for 95% confidence level.
- 22 AUBERT 03P measures  $Y \equiv 2 \tau^0 / (\tau^+ + \tau^-) - 1$ , where  $\tau^0$  is the  $D^0 \rightarrow K^- \pi^+$  (and  $\overline{D}^0 \rightarrow K^+ \pi^-$ ) lifetime, and  $\tau^+$  and  $\tau^-$  are the  $D^0$  and  $\overline{D}^0$  lifetimes to  $CP$ -even states (here  $K^- K^+$  and  $\pi^- \pi^+$ ). In the limit of  $CP$  conservation,  $Y = y \equiv \Delta\Gamma / 2\Gamma$  (we list  $2y = \Delta\Gamma/\Gamma$ ). AUBERT 03P also uses  $\tau^+ - \tau^-$  to get  $\Delta Y = -0.008 \pm 0.006 \pm 0.002$ .

 $|q/p|$ 

The mass eigenstates  $D_1^0$  and  $D_2^0$  are related to the  $C = \pm 1$  states by  $|D_{1,2}^0\rangle = p|D^0\rangle + q|\overline{D}^0\rangle$ . See the note on " $D^0 - \overline{D}^0$  Mixing" above.

"OUR EVALUATION" comes from CPV allowing averages provided by the Heavy Flavor Averaging Group. This would include as-yet-unpublished results, see the note on " $D^0 - \overline{D}^0$  Mixing."

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.92<math>^{+0.12}_{-0.09}</math> OUR EVALUATION</b>	HFAG fit; see the note on " $D^0 - \overline{D}^0$ Mixing."		

<b>0.90<math>^{+0.16+0.08}_{-0.15-0.06}</math></b>	1 PENG	14 BELL	$e^+ e^- \rightarrow \gamma(nS)$
	2 AAIJ	13CE LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.86 $^{+0.30+0.10}_{-0.29-0.08}$	3 ZHANG	07B BELL	Repl. by PENG 14

- 1 The time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$  is employed. Decay-time information and interference on the Dalitz plot are used to distinguish doubly Cabibbo-suppressed decays from mixing and to measure the relative phase between  $D^0 \rightarrow K^{*+} \pi^-$  and  $\overline{D}^0 \rightarrow K^{*+} \pi^-$ . This value allows  $CP$  violation and is sensitive to the sign of  $\Delta m$ .
- 2 Based on 3 fb $^{-1}$  of data collected at  $\sqrt{s} = 7, 8$  TeV. Allowing for  $CP$  violation, the direct  $CP$  violation in mixing is reported  $0.75 < |q/p| < 1.24$  at the 68.3% CL for the  $D^0 \rightarrow K^+ \pi^-$  and  $\overline{D}^0 \rightarrow K^+ \pi^-$ .
- 3 The phase of  $p/q$  is  $(-14 \pm 16 \pm 5)^\circ$ . The ZHANG 07B value is from the time-dependent Dalitz-plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$ . Decay-time information and interference on the Dalitz plot are used to distinguish doubly Cabibbo-suppressed decays from mixing and to measure the relative phase between  $D^0 \rightarrow K^{*+} \pi^-$  and  $\overline{D}^0 \rightarrow K^{*+} \pi^-$ . This value allows  $CP$  violation.

 $A_\Gamma$ 

$A_\Gamma$  is the decay-rate asymmetry for  $CP$ -even final states  $A_\Gamma = (\overline{\tau}_+ - \tau_+) / (\overline{\tau}_+ + \tau_+)$ . See the note on " $D^0 - \overline{D}^0$  Mixing" above.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.125 <math>\pm</math> 0.526 OUR EVALUATION</b>				
<b>-0.6 <math>\pm</math> 0.4 OUR AVERAGE</b>				
-0.3 $\pm$ 2.0 $\pm$ 0.7	1	STARIC	16	BELL $e^+ e^- \rightarrow \gamma(nS)$
-1.34 $\pm$ 0.77 $\pm$ 0.26 -0.34	2	AAIJ	15AA	LHCB $pp$ at 7, 8 TeV
-0.92 $\pm$ 1.45 $\pm$ 0.25 -0.33	3	AAIJ	15AA	LHCB $pp$ at 7, 8 TeV
-0.35 $\pm$ 0.62 $\pm$ 0.12	4	AAIJ	14AL	LHCB $pp$ at 7 TeV
0.33 $\pm$ 1.06 $\pm$ 0.14	5	AAIJ	14AL	LHCB $pp$ at 7 TeV
-1.2 $\pm$ 1.2	6	AALTONEN	14Q	CDF $p\overline{p}$ , $\sqrt{s} = 1.96$ TeV
0.9 $\pm$ 2.6 $\pm$ 0.6		LEES	13	BABR $e^+ e^- \rightarrow \gamma(4S)$
-5.9 $\pm$ 5.9 $\pm$ 2.1	4	AAIJ	12K	LHCB $pp$ at 7 TeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
2.6 $\pm$ 3.6 $\pm$ 0.8		AUBERT	08U	BABR See LEES 13
0.1 $\pm$ 3.0 $\pm$ 2.5		STARIC	07	BELL Repl. by STARIC 16
8 $\pm$ 6 $\pm$ 2		AUBERT	03P	BELL $e^+ e^- \approx \gamma(4S)$

- 1 An improved measurement of  $\overline{D}^0 - D^0$  mixing and a search for  $CP$  violation in  $D^0$  decays to  $CP$ -even final states  $K^+ K^-$  and  $\pi^+ \pi^-$  using the final Belle data sample of 976 fb $^{-1}$ .
- 2 Measured using  $D^0 \rightarrow K^+ K^-$  decays, with  $D^0$  from partially reconstructed semileptonic  $B$  hadron decays.
- 3 Measured using  $D^0 \rightarrow \pi^+ \pi^-$  decays, with  $D^0$  from partially reconstructed semileptonic  $B$  hadron decays.
- 4 Measured using  $D^{*+} \rightarrow D^0 \pi^+$ ,  $D^0 \rightarrow K^+ K^-$  decays (and cc).
- 5 Measured using  $D^{*+} \rightarrow D^0 \pi^+$ ,  $D^0 \rightarrow \pi^+ \pi^-$  decays (and cc).
- 6 Combined result from  $D^0 \rightarrow K^+ K^-$  and  $D^0 \rightarrow \pi^+ \pi^-$ , with  $D^0$  from  $D^{*+} \rightarrow D^0 \pi^+$  (and cc).

 $\cos \delta$ 

$\delta$  is the  $D^0 \rightarrow K^+ \pi^-$  relative strong phase.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.97<math>\pm 0.11</math> OUR AVERAGE</b>			
1.02 $\pm 0.11 \pm 0.06$	1 ABLIKIM	14c BES3	$e^+ e^- \rightarrow D^0 \overline{D}^0$ , 3.77 GeV
0.81 $\pm 0.22 \pm 0.07_{-0.18-0.05}$	2 ASNER	12 CLEO	$e^+ e^- \rightarrow D^0 \overline{D}^0$ , 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.03 $^{+0.31}_{-0.17} \pm 0.06$	3 ASNER	08 CLEO	Repl. by ASNER 12

# Meson Particle Listings

## D<sup>0</sup>

<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$  to measure the asymmetry of the branching fraction of  $D^0 \rightarrow K^-\pi^+$  in  $CP$ -odd and  $CP$ -even eigenstates to be  $(12.7 \pm 1.3 \pm 0.7)\%$ . A fit that includes external measurements of charm mixing parameters finds the value quoted above.

<sup>2</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where decay rates of  $CP$ -tagged  $K\pi$  final states depend on the strong phases between the decays of  $D^0 \rightarrow K^+\pi^-$  and  $\bar{D}^0 \rightarrow K^+\pi^-$ . The measurements obtained  $\sin(\delta) = -0.01 \pm 0.41 \pm 0.04$  and  $|\delta| = (10^{+28+13}_{-53-00})^\circ$  as well. A fit that includes external measurements of charm mixing parameters finds  $\cos(\delta) = 1.15^{+0.19+0.00}_{-0.17-0.08}$ ,  $\sin(\delta) = 0.56^{+0.32+0.21}_{-0.31-0.20}$ , and  $|\delta| = (18^{+11}_{-17})^\circ$ .

<sup>3</sup> ASNER 08 uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where decay rates of  $CP$ -tagged  $K\pi$  final states depend on  $\cos\delta$  because of interfering amplitudes. The above measurement implies  $|\delta| < 75^\circ$  with a confidence level of 95%. A fit that includes external measurements of charm mixing parameters finds  $\cos\delta = 1.10 \pm 0.35 \pm 0.07$ . See also the note on “ $D^0$ - $\bar{D}^0$  Mixing” p. 783 in our 2008 Review (PDG 08).

### D<sup>0</sup> → K<sup>-</sup>π<sup>+</sup>π<sup>0</sup> COHERENCE FACTOR R<sub>Kππ<sup>0</sup></sub>

See the note on ‘ $D^0$ - $\bar{D}^0$  Mixing’ for the definition.  $R_{K\pi\pi^0}$  can have any value between 0 and 1. A value near 1 indicates the decay is dominated by a few intermediate states with limited interference.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.82±0.07</b>	<sup>1</sup> LIBBY	14	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at $\psi(3770)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	

0.78 <sup>+0.11</sup> <sub>-0.25</sub>	<sup>2</sup> LOWREY	09	CLEO Repl. by LIBBY 14
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<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^+\pi^0$  final states depend on  $R_{K\pi\pi^0}$  and  $\delta K\pi\pi^0$ .

<sup>2</sup> LOWREY 09 uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^+\pi^0$  final states depend on  $R_{K\pi\pi^0}$  and  $\delta K\pi\pi^0$ . A fit that includes external measurements of charm mixing parameters gets  $R_{K\pi\pi^0} = 0.84 \pm 0.07$ .

### D<sup>0</sup> → K<sup>-</sup>π<sup>+</sup>π<sup>0</sup> AVERAGE RELATIVE STRONG PHASE δ<sub>Kππ<sup>0</sup></sub>

The quoted value of  $\delta$  is based on the same sign  $CP$  phase of  $D^0$  and  $\bar{D}^0$  convention.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>164<sup>+20</sup><sub>-14</sub></b>	<sup>1</sup> LIBBY	14	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at $\psi(3770)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	

239 <sup>+32</sup> <sub>-28</sub>	<sup>2</sup> LOWREY	09	CLEO Repl. by LIBBY 14
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<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^+\pi^0$  final states depend on  $R_{K\pi\pi^0}$  and  $\delta K\pi\pi^0$ .

<sup>2</sup> LOWREY 09 uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^+\pi^0$  final states depend on  $R_{K\pi\pi^0}$  and  $\delta K\pi\pi^0$ .

A fit that includes external measurements of charm mixing parameters gets  $\delta K\pi\pi^0 = (227^{+14}_{-17})^\circ$ .

### D<sup>0</sup> → K<sup>-</sup>π<sup>-</sup>2π<sup>+</sup> COHERENCE FACTOR R<sub>K3π</sub>

See the note on ‘ $D^0$ - $\bar{D}^0$  Mixing’ for the definition.  $R_{K3\pi}$  can have any value between 0 and 1. A value near 1 indicates the decay is dominated by a few intermediate states with limited interference.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.32<sup>+0.20</sup><sub>-0.28</sub></b>	<sup>1</sup> LIBBY	14	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at $\psi(3770)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	

0.36 <sup>+0.24</sup> <sub>-0.30</sub>	<sup>2</sup> LOWREY	09	CLEO Repl. by LIBBY 14
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<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^-\pi^+\pi^0$  final states depend on  $R_{K3\pi}$  and  $\delta K3\pi$ .

<sup>2</sup> LOWREY 09 uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^-\pi^+\pi^0$  final states depend on  $R_{K3\pi}$  and  $\delta K3\pi$ . A fit that includes external measurements of charm mixing parameters gets  $R_{K3\pi} = 0.33^{+0.26}_{-0.23}$ .

### D<sup>0</sup> → K<sup>-</sup>π<sup>-</sup>2π<sup>+</sup> AVERAGE RELATIVE STRONG PHASE δ<sub>K3π</sub>

The quoted value of  $\delta$  is based on the same sign  $CP$  phase of  $D^0$  and  $\bar{D}^0$  convention.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>225<sup>+21</sup><sub>-78</sub></b>	<sup>1</sup> LIBBY	14	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at $\psi(3770)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	

118 <sup>+62</sup> <sub>-53</sub>	<sup>2</sup> LOWREY	09	CLEO Repl. by LIBBY 14
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<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^-\pi^+\pi^0$  final states depend on  $R_{K3\pi}$  and  $\delta K3\pi$ .

<sup>2</sup> LOWREY 09 uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the decay rates of  $CP$ -tagged  $K^-\pi^-\pi^+\pi^0$  final states depend on  $R_{K3\pi}$  and  $\delta K3\pi$ . A fit that includes external measurements of charm mixing parameters gets  $\delta K3\pi = (114^{+29}_{-23})^\circ$ .

### D<sup>0</sup> → K<sup>0</sup><sub>S</sub>K<sup>+</sup>π<sup>-</sup> COHERENCE FACTOR R<sub>K<sup>0</sup><sub>S</sub>Kπ</sub>

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.73±0.08</b>	<sup>1</sup> INSLER	12	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at 3.77 GeV

<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the signal side  $D$  decays to  $K^0_S K\pi$  and the tag-side  $D$  decays to  $K\pi$ ,  $K\pi\pi\pi$ ,  $K\pi\pi^0$ .

### D<sup>0</sup> → K<sup>0</sup><sub>S</sub>K<sup>+</sup>π<sup>-</sup> AVERAGE RELATIVE STRONG PHASE δ<sub>K<sup>0</sup><sub>S</sub>Kπ</sub>

The quoted value of  $\delta$  is based on the same sign  $CP$  phase of  $D^0$  and  $\bar{D}^0$  convention.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>8.3±15.2</b>	<sup>1</sup> INSLER	12	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at 3.77 GeV

<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the signal side  $D$  decays to  $K^0_S K\pi$  and the tag-side  $D$  decays to  $K\pi$ ,  $K\pi\pi\pi$ ,  $K\pi\pi^0$ .

### D<sup>0</sup> → K<sup>\*</sup>K COHERENCE FACTOR R<sub>K<sup>\*</sup>K</sub>

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.00±0.16</b>	<sup>1</sup> INSLER	12	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at 3.77 GeV

<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the signal side  $D$  decays to  $K^0_S K\pi$  and the tag-side  $D$  decays to  $K\pi$ ,  $K\pi\pi\pi$ ,  $K\pi\pi^0$ .

### D<sup>0</sup> → K<sup>\*</sup>K AVERAGE RELATIVE STRONG PHASE δ<sub>K<sup>\*</sup>K</sub>

The quoted value of  $\delta$  is based on the same sign  $CP$  phase of  $D^0$  and  $\bar{D}^0$  convention.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>26.5±15.8</b>	<sup>1</sup> INSLER	12	CLEO $e^+e^- \rightarrow D^0\bar{D}^0$ at 3.77 GeV

<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the signal side  $D$  decays to  $K^0_S K\pi$  and the tag-side  $D$  decays to  $K\pi$ ,  $K\pi\pi\pi$ ,  $K\pi\pi^0$ .

## D<sup>0</sup> DECAY MODES

Most decay modes (other than the semileptonic modes) that involve a neutral  $K$  meson are now given as  $K^0_S$  modes, not as  $\bar{K}^0$  modes. Nearly always it is a  $K^0_S$  that is measured, and interference between Cabibbo-allowed and doubly Cabibbo-suppressed modes can invalidate the assumption that  $2\Gamma(K^0_S) = \Gamma(\bar{K}^0)$ .

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
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### Topological modes

Γ <sub>1</sub>	0-prongs	[a] (15 ± 6 ) %
Γ <sub>2</sub>	2-prongs	(70 ± 6 ) %
Γ <sub>3</sub>	4-prongs	[b] (14.5 ± 0.5 ) %
Γ <sub>4</sub>	6-prongs	[c] ( 6.4 ± 1.3 ) × 10 <sup>-4</sup>

### Inclusive modes

Γ <sub>5</sub>	$e^+$ anything	[d] ( 6.49 ± 0.11 ) %	
Γ <sub>6</sub>	$\mu^+$ anything	( 6.7 ± 0.6 ) %	
Γ <sub>7</sub>	$K^+$ anything	(54.7 ± 2.8 ) %	S=1.3
Γ <sub>8</sub>	$\bar{K}^0$ anything + $K^0$ anything	(47 ± 4 ) %	
Γ <sub>9</sub>	$K^+$ anything	( 3.4 ± 0.4 ) %	
Γ <sub>10</sub>	$K^*(892)^-$ anything	(15 ± 9 ) %	
Γ <sub>11</sub>	$\bar{K}^*(892)^0$ anything	( 9 ± 4 ) %	
Γ <sub>12</sub>	$K^*(892)^+$ anything	< 3.6 %	CL=90%
Γ <sub>13</sub>	$K^*(892)^0$ anything	( 2.8 ± 1.3 ) %	
Γ <sub>14</sub>	$\eta$ anything	( 9.5 ± 0.9 ) %	
Γ <sub>15</sub>	$\eta'$ anything	( 2.48 ± 0.27 ) %	
Γ <sub>16</sub>	$\phi$ anything	( 1.05 ± 0.11 ) %	

### Semileptonic modes

Γ <sub>17</sub>	$K^-\ell^+\nu_\ell$		
Γ <sub>18</sub>	$K^+e^+\nu_e$	( 3.538 ± 0.033 ) %	S=1.3
Γ <sub>19</sub>	$K^-\mu^+\nu_\mu$	( 3.33 ± 0.13 ) %	
Γ <sub>20</sub>	$K^*(892)^-e^+\nu_e$	( 2.16 ± 0.16 ) %	
Γ <sub>21</sub>	$K^*(892)^-\mu^+\nu_\mu$	( 1.92 ± 0.25 ) %	
Γ <sub>22</sub>	$K^-\pi^0e^+\nu_e$	( 1.6 <sup>+1.3</sup> <sub>-0.5</sub> ) %	
Γ <sub>23</sub>	$\bar{K}^0\pi^-e^+\nu_e$	( 2.7 <sup>+0.9</sup> <sub>-0.7</sub> ) %	
Γ <sub>24</sub>	$K^-\pi^+\pi^0e^+\nu_e$	( 2.8 <sup>+1.4</sup> <sub>-1.1</sub> ) × 10 <sup>-4</sup>	
Γ <sub>25</sub>	$K_1(1270)^-e^+\nu_e$	( 7.6 <sup>+4.0</sup> <sub>-3.1</sub> ) × 10 <sup>-4</sup>	
Γ <sub>26</sub>	$K^-\pi^+\pi^-\mu^+\nu_\mu$	< 1.2 × 10 <sup>-3</sup>	CL=90%
Γ <sub>27</sub>	$(\bar{K}^*(892)\pi)^-\mu^+\nu_\mu$	< 1.4 × 10 <sup>-3</sup>	CL=90%
Γ <sub>28</sub>	$\pi^+e^+\nu_e$	( 2.91 ± 0.04 ) × 10 <sup>-3</sup>	S=1.1
Γ <sub>29</sub>	$\pi^-\mu^+\nu_\mu$	( 2.38 ± 0.24 ) × 10 <sup>-3</sup>	
Γ <sub>30</sub>	$\rho^-e^+\nu_e$	( 1.77 ± 0.16 ) × 10 <sup>-3</sup>	

Hadronic modes with one $\bar{K}$							
$\Gamma_{31}$	$K^- \pi^+$	( 3.93 $\pm$ 0.04 ) %	S=1.2	$\Gamma_{76}$	$K_S^0 \pi^+ \pi^- \pi^0$	[h] ( 5.2 $\pm$ 0.6 ) %	
$\Gamma_{32}$	$K^+ \pi^-$	( 1.398 $\pm$ 0.027 ) $\times 10^{-4}$		$\Gamma_{77}$	$K_S^0 \eta, \eta \rightarrow \pi^+ \pi^- \pi^0$	( 1.02 $\pm$ 0.09 ) $\times 10^{-3}$	
$\Gamma_{33}$	$K_S^0 \pi^0$	( 1.20 $\pm$ 0.04 ) %		$\Gamma_{78}$	$K_S^0 \omega, \omega \rightarrow \pi^+ \pi^- \pi^0$	( 9.9 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{34}$	$K_S^0 \pi^0$	( 10.0 $\pm$ 0.7 ) $\times 10^{-3}$		$\Gamma_{79}$	$K^- \pi^+ 2\pi^0$		
$\Gamma_{35}$	$K_S^0 \pi^+ \pi^-$	[e] ( 2.85 $\pm$ 0.20 ) %	S=1.1	$\Gamma_{80}$	$K^- 2\pi^+ \pi^- \pi^0$	( 4.2 $\pm$ 0.4 ) %	
$\Gamma_{36}$	$K_S^0 \rho^0$	( 6.4 $\pm$ 0.7 ) $\times 10^{-3}$		$\Gamma_{81}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.3 $\pm$ 0.6 ) %	
$\Gamma_{37}$	$K_S^0 \omega, \omega \rightarrow \pi^+ \pi^-$	( 2.1 $\pm$ 0.6 ) $\times 10^{-4}$		$\Gamma_{82}$	$K^- \pi^+ \omega, \omega \rightarrow \pi^+ \pi^- \pi^0$	( 2.7 $\pm$ 0.5 ) %	
$\Gamma_{38}$	$K_S^0 (\pi^+ \pi^-)_{S\text{-wave}}$	( 3.4 $\pm$ 0.8 ) $\times 10^{-3}$		$\Gamma_{83}$	$\bar{K}^*(892)^0 \omega,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+,$ $\omega \rightarrow \pi^+ \pi^- \pi^0$	( 6.5 $\pm$ 3.0 ) $\times 10^{-3}$	
$\Gamma_{39}$	$K_S^0 f_0(980),$ $f_0(980) \rightarrow \pi^+ \pi^-$	( 1.23 $\pm$ 0.40 ) $\times 10^{-3}$		$\Gamma_{84}$	$K_S^0 \eta \pi^0$	( 5.5 $\pm$ 1.1 ) $\times 10^{-3}$	
$\Gamma_{40}$	$K_S^0 f_0(1370),$ $f_0(1370) \rightarrow \pi^+ \pi^-$	( 2.8 $\pm$ 0.9 ) $\times 10^{-3}$		$\Gamma_{85}$	$K_S^0 a_0(980), a_0(980) \rightarrow \eta \pi^0$	( 6.6 $\pm$ 2.0 ) $\times 10^{-3}$	
$\Gamma_{41}$	$K_S^0 f_2(1270),$ $f_2(1270) \rightarrow \pi^+ \pi^-$	( 9 $\pm$ 10 ) $\times 10^{-5}$		$\Gamma_{86}$	$\bar{K}^*(892)^0 \eta, \bar{K}^*(892)^0 \rightarrow$ $K_S^0 \pi^0$	( 1.6 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{42}$	$K^*(892)^- \pi^+,$ $K^*(892)^- \rightarrow K_S^0 \pi^-$	( 1.68 $\pm$ 0.15 ) %		$\Gamma_{87}$	$K_S^0 2\pi^+ 2\pi^-$	( 2.71 $\pm$ 0.31 ) $\times 10^{-3}$	
$\Gamma_{43}$	$K_0^*(1430)^- \pi^+,$ $K_0^*(1430)^- \rightarrow K_S^0 \pi^-$	( 2.73 $\pm$ 0.40 ) $\times 10^{-3}$		$\Gamma_{88}$	$K_S^0 \rho^0 \pi^+ \pi^-, \text{ no } K^*(892)^-$	( 1.1 $\pm$ 0.7 ) $\times 10^{-3}$	
$\Gamma_{44}$	$K_2^*(1430)^- \pi^+,$ $K_2^*(1430)^- \rightarrow K_S^0 \pi^-$	( 3.4 $\pm$ 1.9 ) $\times 10^{-4}$		$\Gamma_{89}$	$K^*(892)^- 2\pi^+ \pi^-,$ $K^*(892)^- \rightarrow K_S^0 \pi^-, \text{ no}$ $\rho^0$	( 5 $\pm$ 8 ) $\times 10^{-4}$	
$\Gamma_{45}$	$K^*(1680)^- \pi^+,$ $K^*(1680)^- \rightarrow K_S^0 \pi^-$	( 4 $\pm$ 4 ) $\times 10^{-4}$		$\Gamma_{90}$	$K^*(892)^- \rho^0 \pi^+,$ $K^*(892)^- \rightarrow K_S^0 \pi^-$	( 1.6 $\pm$ 0.6 ) $\times 10^{-3}$	
$\Gamma_{46}$	$K^*(892)^+ \pi^-,$ $K^*(892)^+ \rightarrow K_S^0 \pi^+$	[f] ( 1.15 $\pm$ 0.60 ) $\times 10^{-4}$		$\Gamma_{91}$	$K_S^0 2\pi^+ 2\pi^- \text{ nonresonant}$	< 1.2 $\times 10^{-3}$	CL=90%
$\Gamma_{47}$	$K_0^*(1430)^+ \pi^-,$ $K_0^*(1430)^+ \rightarrow K_S^0 \pi^+$	[f] < 1.4 $\times 10^{-5}$	CL=95%	$\Gamma_{92}$	$\bar{K}^0 \pi^+ \pi^- 2\pi^0 (\pi^0)$		
$\Gamma_{48}$	$K_2^*(1430)^+ \pi^-,$ $K_2^*(1430)^+ \rightarrow K_S^0 \pi^+$	[f] < 3.4 $\times 10^{-5}$	CL=95%	$\Gamma_{93}$	$K^- 3\pi^+ 2\pi^-$	( 2.2 $\pm$ 0.6 ) $\times 10^{-4}$	
$\Gamma_{49}$	$K_S^0 \pi^+ \pi^- \text{ nonresonant}$	( 2.6 $\pm$ 6.0 ) $\times 10^{-4}$		Fractions of many of the following modes with resonances have already appeared above as submodes of particular charged-particle modes. (Modes for which there are only upper limits and $\bar{K}^*(892)^0 \rho$ submodes only appear below.)			
$\Gamma_{50}$	$K^- \pi^+ \pi^0$	[e] ( 14.3 $\pm$ 0.8 ) %	S=3.1	$\Gamma_{94}$	$K_S^0 \eta$	( 4.85 $\pm$ 0.30 ) $\times 10^{-3}$	
$\Gamma_{51}$	$K^- \rho^+$	( 11.1 $\pm$ 0.9 ) %		$\Gamma_{95}$	$K_S^0 \omega$	( 1.11 $\pm$ 0.06 ) %	
$\Gamma_{52}$	$K^- \rho(1700)^+,$ $\rho(1700)^+ \rightarrow \pi^+ \pi^0$	( 8.1 $\pm$ 1.8 ) $\times 10^{-3}$		$\Gamma_{96}$	$K_S^0 \eta'(958)$	( 9.5 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{53}$	$K^*(892)^- \pi^+,$ $K^*(892)^- \rightarrow K^- \pi^0$	( 2.28 $\pm$ 0.40 ) %		$\Gamma_{97}$	$K^- a_1(1260)^+$	( 7.8 $\pm$ 1.1 ) %	
$\Gamma_{54}$	$\bar{K}^*(892)^0 \pi^0,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.93 $\pm$ 0.26 ) %		$\Gamma_{98}$	$K^- a_2(1320)^+$	< 2 $\times 10^{-3}$	CL=90%
$\Gamma_{55}$	$K_0^*(1430)^- \pi^+,$ $K_0^*(1430)^- \rightarrow K^- \pi^0$	( 4.7 $\pm$ 2.2 ) $\times 10^{-3}$		$\Gamma_{99}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \text{ total}$	( 2.4 $\pm$ 0.5 ) %	
$\Gamma_{56}$	$\bar{K}_0^*(1430)^0 \pi^0,$ $\bar{K}_0^*(1430)^0 \rightarrow K^- \pi^+$	( 5.8 $\pm$ 5.0 ) $\times 10^{-3}$		$\Gamma_{100}$	$\bar{K}^*(892)^0 \pi^+ \pi^- 3\text{-body}$	( 1.48 $\pm$ 0.34 ) %	
$\Gamma_{57}$	$K^*(1680)^- \pi^+,$ $K^*(1680)^- \rightarrow K^- \pi^0$	( 1.9 $\pm$ 0.7 ) $\times 10^{-3}$		$\Gamma_{101}$	$\bar{K}^*(892)^0 \rho^0$	( 1.57 $\pm$ 0.35 ) %	
$\Gamma_{58}$	$K^- \pi^+ \pi^0 \text{ nonresonant}$	( 1.14 $\pm$ 0.50 ) %		$\Gamma_{102}$	$\bar{K}^*(892)^0 \rho^0 \text{ transverse}$	( 1.7 $\pm$ 0.6 ) %	
$\Gamma_{59}$	$K_S^0 2\pi^0$	( 9.1 $\pm$ 1.1 ) $\times 10^{-3}$	S=2.2	$\Gamma_{103}$	$\bar{K}^*(892)^0 \rho^0 S\text{-wave}$	( 3.0 $\pm$ 0.6 ) %	
$\Gamma_{60}$	$K_S^0 (2\pi^0)\text{-}S\text{-wave}$	( 2.6 $\pm$ 0.7 ) $\times 10^{-3}$		$\Gamma_{104}$	$\bar{K}^*(892)^0 \rho^0 S\text{-wave long.}$	< 3 $\times 10^{-3}$	CL=90%
$\Gamma_{61}$	$\bar{K}^*(892)^0 \pi^0,$ $\bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0$	( 7.9 $\pm$ 0.7 ) $\times 10^{-3}$		$\Gamma_{105}$	$\bar{K}^*(892)^0 \rho^0 P\text{-wave}$	< 3 $\times 10^{-3}$	CL=90%
$\Gamma_{62}$	$\bar{K}^*(1430)^0 \pi^0, \bar{K}^{*0} \rightarrow K_S^0 \pi^0$	( 4 $\pm$ 23 ) $\times 10^{-5}$		$\Gamma_{106}$	$\bar{K}^*(892)^0 \rho^0 D\text{-wave}$	( 2.1 $\pm$ 0.6 ) %	
$\Gamma_{63}$	$K^*(1680)^0 \pi^0, \bar{K}^{*0} \rightarrow K_S^0 \pi^0$	( 1.0 $\pm$ 0.4 ) $\times 10^{-3}$		$\Gamma_{107}$	$K^- \pi^+ f_0(980)$		
$\Gamma_{64}$	$K_S^0 f_2(1270), f_2 \rightarrow 2\pi^0$	( 2.3 $\pm$ 1.1 ) $\times 10^{-4}$		$\Gamma_{108}$	$\bar{K}^*(892)^0 f_0(980)$		
$\Gamma_{65}$	$2K_S^0, \text{ one } K_S^0 \rightarrow 2\pi^0$	( 3.2 $\pm$ 1.1 ) $\times 10^{-4}$		$\Gamma_{109}$	$K_1(1270)^- \pi^+$	[g] ( 1.6 $\pm$ 0.8 ) %	
$\Gamma_{66}$	$K_S^0 2\pi^0 \text{ nonresonant}$			$\Gamma_{110}$	$K_1(1400)^- \pi^+$	< 1.2 %	CL=90%
$\Gamma_{67}$	$K^- 2\pi^+ \pi^-$	[e] ( 8.06 $\pm$ 0.23 ) %	S=1.5	$\Gamma_{111}$	$K^*(1410)^- \pi^+$		
$\Gamma_{68}$	$K^- \pi^+ \rho^0 \text{ total}$	( 6.73 $\pm$ 0.34 ) %		$\Gamma_{112}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0$	( 1.9 $\pm$ 0.9 ) %	
$\Gamma_{69}$	$K^- \pi^+ \rho^0 3\text{-body}$	( 5.1 $\pm$ 2.3 ) $\times 10^{-3}$		$\Gamma_{113}$	$\bar{K}^*(892)^0 \eta$		
$\Gamma_{70}$	$\bar{K}^*(892)^0 \rho^0,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.05 $\pm$ 0.23 ) %		$\Gamma_{114}$	$K^- \pi^+ \omega$	( 3.1 $\pm$ 0.6 ) %	
$\Gamma_{71}$	$K^- a_1(1260)^+,$ $a_1(1260)^+ \rightarrow 2\pi^+ \pi^-$	( 3.6 $\pm$ 0.6 ) %		$\Gamma_{115}$	$\bar{K}^*(892)^0 \omega$	( 1.1 $\pm$ 0.5 ) %	
$\Gamma_{72}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \text{ total},$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.6 $\pm$ 0.4 ) %		$\Gamma_{116}$	$K^- \pi^+ \eta'(958)$	( 7.5 $\pm$ 1.9 ) $\times 10^{-3}$	
$\Gamma_{73}$	$\bar{K}^*(892)^0 \pi^+ \pi^- 3\text{-body},$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 9.9 $\pm$ 2.3 ) $\times 10^{-3}$		$\Gamma_{117}$	$\bar{K}^*(892)^0 \eta'(958)$	< 1.1 $\times 10^{-3}$	CL=90%
$\Gamma_{74}$	$K_1(1270)^- \pi^+,$ $K_1(1270)^- \rightarrow K^- \pi^+ \pi^-$	[g] ( 2.9 $\pm$ 0.3 ) $\times 10^{-3}$		Hadronic modes with three $K$ 's			
$\Gamma_{75}$	$K^- 2\pi^+ \pi^- \text{ nonresonant}$	( 1.88 $\pm$ 0.26 ) %		$\Gamma_{118}$	$K_S^0 K^+ K^-$	( 4.51 $\pm$ 0.34 ) $\times 10^{-3}$	
				$\Gamma_{119}$	$K_S^0 a_0(980)^0, a_0^0 \rightarrow K^+ K^-$	( 3.0 $\pm$ 0.4 ) $\times 10^{-3}$	
				$\Gamma_{120}$	$K^- a_0(980)^+, a_0^+ \rightarrow K^+ K_S^0$	( 6.0 $\pm$ 1.8 ) $\times 10^{-4}$	
				$\Gamma_{121}$	$K^+ a_0(980)^-, a_0^- \rightarrow K^- K_S^0$	< 1.1 $\times 10^{-4}$	CL=95%
				$\Gamma_{122}$	$K_S^0 f_0(980), f_0 \rightarrow K^+ K^-$	< 9 $\times 10^{-5}$	CL=95%
				$\Gamma_{123}$	$K_S^0 \phi, \phi \rightarrow K^+ K^-$	( 2.07 $\pm$ 0.16 ) $\times 10^{-3}$	
				$\Gamma_{124}$	$K_S^0 f_0(1370), f_0 \rightarrow K^+ K^-$	( 1.7 $\pm$ 1.1 ) $\times 10^{-4}$	
				$\Gamma_{125}$	$3K_S^0$	( 9.2 $\pm$ 1.3 ) $\times 10^{-4}$	
				$\Gamma_{126}$	$K^+ 2K^- \pi^+$	( 2.21 $\pm$ 0.32 ) $\times 10^{-4}$	
				$\Gamma_{127}$	$K^+ K^- \bar{K}^*(892)^0,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 4.4 $\pm$ 1.7 ) $\times 10^{-5}$	
				$\Gamma_{128}$	$K^- \pi^+ \phi, \phi \rightarrow K^+ K^-$	( 4.0 $\pm$ 1.7 ) $\times 10^{-5}$	
				$\Gamma_{129}$	$\phi \bar{K}^*(892)^0,$ $\phi \rightarrow K^+ K^-,$ $\bar{K}^*(892)^0 \rightarrow K^- \pi^+$	( 1.06 $\pm$ 0.20 ) $\times 10^{-4}$	
				$\Gamma_{130}$	$K^+ 2K^- \pi^+ \text{ nonresonant}$	( 3.3 $\pm$ 1.5 ) $\times 10^{-5}$	
				$\Gamma_{131}$	$2K_S^0 K^\pm \pi^\mp$	( 6.1 $\pm$ 1.3 ) $\times 10^{-4}$	

## Meson Particle Listings

 $D^0$ 

Pionic modes			
$\Gamma_{132}$	$\pi^+\pi^-$	$(1.420 \pm 0.025) \times 10^{-3}$	S=1.1
$\Gamma_{133}$	$2\pi^0$	$(8.25 \pm 0.25) \times 10^{-4}$	
$\Gamma_{134}$	$\pi^+\pi^-\pi^0$	$(1.47 \pm 0.09) \%$	S=3.0
$\Gamma_{135}$	$\rho^+\pi^-$	$(1.00 \pm 0.06) \%$	
$\Gamma_{136}$	$\rho^0\pi^0$	$(3.82 \pm 0.29) \times 10^{-3}$	
$\Gamma_{137}$	$\rho^-\pi^+$	$(5.09 \pm 0.34) \times 10^{-3}$	
$\Gamma_{138}$	$\rho(1450)^+\pi^-, \rho(1450)^+ \rightarrow \pi^+\pi^-\pi^0$	$(1.6 \pm 2.0) \times 10^{-5}$	
$\Gamma_{139}$	$\rho(1450)^0\pi^0, \rho(1450)^0 \rightarrow \pi^+\pi^-$	$(4.4 \pm 1.9) \times 10^{-5}$	
$\Gamma_{140}$	$\rho(1450)^-\pi^+, \rho(1450)^- \rightarrow \pi^+\pi^-$	$(2.6 \pm 0.4) \times 10^{-4}$	
$\Gamma_{141}$	$\rho(1700)^+\pi^-, \rho(1700)^+ \rightarrow \pi^+\pi^-$	$(6.0 \pm 1.5) \times 10^{-4}$	
$\Gamma_{142}$	$\rho(1700)^0\pi^0, \rho(1700)^0 \rightarrow \pi^+\pi^-$	$(7.4 \pm 1.8) \times 10^{-4}$	
$\Gamma_{143}$	$\rho(1700)^-\pi^+, \rho(1700)^- \rightarrow \pi^+\pi^-$	$(4.7 \pm 1.1) \times 10^{-4}$	
$\Gamma_{144}$	$f_0(980)\pi^0, f_0(980) \rightarrow \pi^+\pi^-$	$(3.7 \pm 0.9) \times 10^{-5}$	
$\Gamma_{145}$	$f_0(500)\pi^0, f_0(500) \rightarrow \pi^+\pi^-$	$(1.21 \pm 0.22) \times 10^{-4}$	
$\Gamma_{146}$	$(\pi^+\pi^-)_{S\text{-wave}}\pi^0$		
$\Gamma_{147}$	$f_0(1370)\pi^0, f_0(1370) \rightarrow \pi^+\pi^-$	$(5.4 \pm 2.1) \times 10^{-5}$	
$\Gamma_{148}$	$f_0(1500)\pi^0, f_0(1500) \rightarrow \pi^+\pi^-$	$(5.7 \pm 1.6) \times 10^{-5}$	
$\Gamma_{149}$	$f_0(1710)\pi^0, f_0(1710) \rightarrow \pi^+\pi^-$	$(4.6 \pm 1.6) \times 10^{-5}$	
$\Gamma_{150}$	$f_2(1270)\pi^0, f_2(1270) \rightarrow \pi^+\pi^-$	$(1.94 \pm 0.22) \times 10^{-4}$	
$\Gamma_{151}$	$\pi^+\pi^-\pi^0$ nonresonant	$(1.2 \pm 0.4) \times 10^{-4}$	
$\Gamma_{152}$	$3\pi^0$	$< 3.5 \times 10^{-4}$	CL=90%
$\Gamma_{153}$	$2\pi^+2\pi^-$	$(7.45 \pm 0.22) \times 10^{-3}$	S=1.2
$\Gamma_{154}$	$a_1(1260)^+\pi^-, a_1^+ \rightarrow 2\pi^+\pi^-$ total	$(4.47 \pm 0.32) \times 10^{-3}$	
$\Gamma_{155}$	$a_1(1260)^+\pi^-, a_1^+ \rightarrow \rho^0\pi^+$ S-wave	$(3.23 \pm 0.25) \times 10^{-3}$	
$\Gamma_{156}$	$a_1(1260)^+\pi^-, a_1^+ \rightarrow \rho^0\pi^+$ D-wave	$(1.9 \pm 0.5) \times 10^{-4}$	
$\Gamma_{157}$	$a_1(1260)^+\pi^-, a_1^+ \rightarrow \sigma\pi^+$	$(6.2 \pm 0.7) \times 10^{-4}$	
$\Gamma_{158}$	$2\rho^0$ total	$(1.83 \pm 0.13) \times 10^{-3}$	
$\Gamma_{159}$	$2\rho^0$ , parallel helicities	$(8.2 \pm 3.2) \times 10^{-5}$	
$\Gamma_{160}$	$2\rho^0$ , perpendicular helicities	$(4.8 \pm 0.6) \times 10^{-4}$	
$\Gamma_{161}$	$2\rho^0$ , longitudinal helicities	$(1.25 \pm 0.10) \times 10^{-3}$	
$\Gamma_{162}$	Resonant $(\pi^+\pi^-)\pi^+\pi^-$ 3-body total	$(1.49 \pm 0.12) \times 10^{-3}$	
$\Gamma_{163}$	$\sigma\pi^+\pi^-$	$(6.1 \pm 0.9) \times 10^{-4}$	
$\Gamma_{164}$	$f_0(980)\pi^+\pi^-, f_0 \rightarrow \pi^+\pi^-$	$(1.8 \pm 0.5) \times 10^{-4}$	
$\Gamma_{165}$	$f_2(1270)\pi^+\pi^-, f_2 \rightarrow \pi^+\pi^-$	$(3.7 \pm 0.6) \times 10^{-4}$	
$\Gamma_{166}$	$\pi^+\pi^-2\pi^0$	$(1.01 \pm 0.09) \%$	
$\Gamma_{167}$	$\eta\pi^0$	$[i] (6.9 \pm 0.7) \times 10^{-4}$	
$\Gamma_{168}$	$\omega\pi^0$	$[i] < 2.6 \times 10^{-4}$	CL=90%
$\Gamma_{169}$	$2\pi^+2\pi^-\pi^0$	$(4.2 \pm 0.5) \times 10^{-3}$	
$\Gamma_{170}$	$\eta\pi^+\pi^-$	$[i] (1.09 \pm 0.16) \times 10^{-3}$	
$\Gamma_{171}$	$\omega\pi^+\pi^-$	$[i] (1.6 \pm 0.5) \times 10^{-3}$	
$\Gamma_{172}$	$3\pi^+3\pi^-$	$(4.2 \pm 1.2) \times 10^{-4}$	
$\Gamma_{173}$	$\eta'(958)\pi^0$	$(9.1 \pm 1.4) \times 10^{-4}$	
$\Gamma_{174}$	$\eta'(958)\pi^+\pi^-$	$(4.5 \pm 1.7) \times 10^{-4}$	
$\Gamma_{175}$	$2\eta$	$(1.70 \pm 0.20) \times 10^{-3}$	
$\Gamma_{176}$	$\eta\eta'(958)$	$(1.06 \pm 0.27) \times 10^{-3}$	
Hadronic modes with a $K\bar{K}$ pair			
$\Gamma_{177}$	$K^+K^-$	$(4.01 \pm 0.07) \times 10^{-3}$	S=1.5
$\Gamma_{178}$	$2K_S^0$	$(1.8 \pm 0.4) \times 10^{-4}$	S=2.5
$\Gamma_{179}$	$K_S^0K^-\pi^+$	$(3.6 \pm 0.5) \times 10^{-3}$	S=1.2
$\Gamma_{180}$	$\bar{K}^*(892)^0K_S^0, \bar{K}^{*0} \rightarrow K^-\pi^+$	$< 5 \times 10^{-4}$	CL=90%
$\Gamma_{181}$	$K_S^0K^+\pi^-$	$(2.2 \pm 0.4) \times 10^{-3}$	S=1.3
$\Gamma_{182}$	$K^*(892)^0K_S^0, K^{*0} \rightarrow K^-\pi^+$	$< 1.8 \times 10^{-4}$	CL=90%
$\Gamma_{183}$	$K^+K^-\pi^0$	$(3.38 \pm 0.21) \times 10^{-3}$	
$\Gamma_{184}$	$K^*(892)^+K^-, K^*(892)^+ \rightarrow K^+\pi^0$	$(1.50 \pm 0.10) \times 10^{-3}$	
$\Gamma_{185}$	$K^*(892)^-K^+, K^*(892)^- \rightarrow K^-\pi^0$	$(5.4 \pm 0.5) \times 10^{-4}$	
$\Gamma_{186}$	$(K^+\pi^0)_{S\text{-wave}}K^-$	$(2.40 \pm 0.21) \times 10^{-3}$	
$\Gamma_{187}$	$(K^-\pi^0)_{S\text{-wave}}K^+$	$(1.3 \pm 0.5) \times 10^{-4}$	
$\Gamma_{188}$	$f_0(980)\pi^0, f_0 \rightarrow K^+K^-$	$(3.5 \pm 0.6) \times 10^{-4}$	
$\Gamma_{189}$	$\phi\pi^0, \phi \rightarrow K^+K^-$	$(6.6 \pm 0.5) \times 10^{-4}$	
$\Gamma_{190}$	$K^+K^-\pi^0$ nonresonant		
$\Gamma_{191}$	$2K_S^0\pi^0$	$< 5.9 \times 10^{-4}$	
$\Gamma_{192}$	$K^+K^-\pi^+\pi^-$	$(2.42 \pm 0.12) \times 10^{-3}$	
$\Gamma_{193}$	$\phi(\pi^+\pi^-)_{S\text{-wave}}, \phi \rightarrow K^+K^-$	$(2.50 \pm 0.34) \times 10^{-4}$	
$\Gamma_{194}$	$(\phi\rho^0)_{S\text{-wave}}, \phi \rightarrow K^+K^-$	$(9.3 \pm 1.2) \times 10^{-4}$	
$\Gamma_{195}$	$(\phi\rho^0)_{D\text{-wave}}, \phi \rightarrow K^+K^-$	$(8.2 \pm 2.3) \times 10^{-5}$	
$\Gamma_{196}$	$(K^{*0}\bar{K}^{*0})_{S\text{-wave}}, K^{*0} \rightarrow K^\pm\pi^\mp$	$(1.48 \pm 0.30) \times 10^{-4}$	
$\Gamma_{197}$	$(K^-\pi^+)_{P\text{-wave}}, (K^+\pi^-)_{S\text{-wave}}, K_1(1270)^+K^-$	$(2.6 \pm 0.5) \times 10^{-4}$	
$\Gamma_{198}$	$K_1(1270)^+K^-$	$(1.8 \pm 0.5) \times 10^{-4}$	
$\Gamma_{199}$	$K_1(1270)^+ \rightarrow K^{*0}\pi^+$	$(1.14 \pm 0.26) \times 10^{-4}$	
$\Gamma_{200}$	$K_1(1270)^+ \rightarrow \rho^0K^+$	$(2.2 \pm 1.2) \times 10^{-5}$	
$\Gamma_{201}$	$K_1(1270)^- \rightarrow \bar{K}^{*0}\pi^-$	$(1.45 \pm 0.25) \times 10^{-4}$	
$\Gamma_{202}$	$K_1(1270)^- \rightarrow \rho^0K^-$	$(1.02 \pm 0.26) \times 10^{-4}$	
$\Gamma_{203}$	$K^*(1410)^+K^-, K^*(1410)^+ \rightarrow K^{*0}\pi^+$	$(1.14 \pm 0.25) \times 10^{-4}$	
$\Gamma_{204}$	$K^*(1410)^-K^+, K^*(1410)^- \rightarrow \bar{K}^{*0}\pi^-$		
$\Gamma_{205}$	$K^+K^-\rho^0$ 3-body		
$\Gamma_{206}$	$f_0(980)\pi^+\pi^-, f_0 \rightarrow K^+K^-$		
$\Gamma_{207}$	$K^*(892)^0K^\mp\pi^\pm$ 3-body, $K^{*0} \rightarrow K^\pm\pi^\mp$	$K^*(892)^0\bar{K}^*(892)^0, K^{*0} \rightarrow K^\pm\pi^\mp$	
$\Gamma_{208}$	$K_1(1270)^\pm K^\mp, K_1(1270)^\pm \rightarrow K^\pm\pi^+\pi^-$		
$\Gamma_{209}$	$K_1(1400)^\pm K^\mp, K_1(1400)^\pm \rightarrow K^\pm\pi^+\pi^-$		
$\Gamma_{210}$	$2K_S^0\pi^+\pi^-$	$(1.24 \pm 0.24) \times 10^{-3}$	
$\Gamma_{211}$	$K_S^0K^-2\pi^+\pi^-$	$< 1.5 \times 10^{-4}$	CL=90%
$\Gamma_{212}$	$K^+K^-\pi^+\pi^-\pi^0$	$(3.1 \pm 2.0) \times 10^{-3}$	
Other $K\bar{K}X$ modes. They include all decay modes of the $\phi, \eta$ , and $\omega$ .			
$\Gamma_{213}$	$\phi\pi^0$		
$\Gamma_{214}$	$\phi\eta$	$(1.4 \pm 0.5) \times 10^{-4}$	
$\Gamma_{215}$	$\phi\omega$	$< 2.1 \times 10^{-3}$	CL=90%
Radiative modes			
$\Gamma_{216}$	$\rho^0\gamma$	$< 2.4 \times 10^{-4}$	CL=90%
$\Gamma_{217}$	$\omega\gamma$	$< 2.4 \times 10^{-4}$	CL=90%
$\Gamma_{218}$	$\phi\gamma$	$(2.73 \pm 0.35) \times 10^{-5}$	
$\Gamma_{219}$	$\bar{K}^*(892)^0\gamma$	$(3.31 \pm 0.34) \times 10^{-4}$	
Doubly Cabibbo suppressed (DC) modes or $\Delta C = 2$ forbidden via mixing (C2M) modes			
$\Gamma_{220}$	$K^+\ell^-\bar{\nu}_\ell$ via $\bar{D}^0$	$< 2.2 \times 10^{-5}$	CL=90%
$\Gamma_{221}$	$K^+$ or $K^*(892)^+e^-\bar{\nu}_e$ via $\bar{D}^0$	$< 6 \times 10^{-5}$	CL=90%
$\Gamma_{222}$	$K^+\pi^-$ DC	$(1.49 \pm 0.07) \times 10^{-4}$	S=2.9
$\Gamma_{223}$	$K^+\pi^-$ via DCS	$(1.33 \pm 0.09) \times 10^{-4}$	
$\Gamma_{224}$	$K^+\pi^-$ via $\bar{D}^0$	$< 1.6 \times 10^{-5}$	CL=95%
$\Gamma_{225}$	$K_S^0\pi^+\pi^-$ in $D^0 \rightarrow \bar{D}^0$	$< 1.8 \times 10^{-4}$	CL=95%
$\Gamma_{226}$	$K^*(892)^+\pi^-, K^*(892)^+ \rightarrow K_S^0\pi^+$ DC	$(1.15 \pm 0.60 \pm 0.34) \times 10^{-4}$	
$\Gamma_{227}$	$K_S^0(1430)^+\pi^-, K_S^0(1430)^+ \rightarrow K_S^0\pi^+$ DC	$< 1.4 \times 10^{-5}$	
$\Gamma_{228}$	$K_S^0(1430)^+\pi^-, K_S^0(1430)^+ \rightarrow K_S^0\pi^+$ DC	$< 3.4 \times 10^{-5}$	
$\Gamma_{229}$	$K^+\pi^-\pi^0$ DC	$(3.13 \pm 0.23) \times 10^{-4}$	
$\Gamma_{230}$	$K^+\pi^-\pi^0$ via $\bar{D}^0$	$(7.5 \pm 0.6) \times 10^{-4}$	
$\Gamma_{231}$	$K^+\pi^+2\pi^-$ DC	$(2.62 \pm 0.11) \times 10^{-4}$	
$\Gamma_{232}$	$K^+\pi^+2\pi^-$ via $\bar{D}^0$	$< 4 \times 10^{-4}$	CL=90%
$\Gamma_{233}$	$K^+\pi^-$ or $K^+\pi^+2\pi^-$ via $\bar{D}^0$		
$\Gamma_{234}$	$\mu^-\text{anything via } \bar{D}^0$	$< 4 \times 10^{-4}$	CL=90%



$\Delta C = 1$  weak neutral current ( $C1$ ) modes,  
Lepton Family number ( $LF$ ) violating modes,  
Lepton ( $L$ ) or Baryon ( $B$ ) number violating modes

	Decay ( $\Gamma$ )	Channel ( $\mathcal{B}$ )	Number of fitting modes		
235	$\gamma\gamma$	$CI$	< 2.2	$\times 10^{-6}$	CL=90%
236	$e^+e^-$	$CI$	< 7.9	$\times 10^{-8}$	CL=90%
237	$\mu^+\mu^-$	$CI$	< 6.2	$\times 10^{-9}$	CL=90%
238	$\pi^0 e^+e^-$	$CI$	< 4.5	$\times 10^{-5}$	CL=90%
239	$\pi^0 \mu^+\mu^-$	$CI$	< 1.8	$\times 10^{-4}$	CL=90%
240	$\eta e^+e^-$	$CI$	< 1.1	$\times 10^{-4}$	CL=90%
241	$\eta \mu^+\mu^-$	$CI$	< 5.3	$\times 10^{-4}$	CL=90%
242	$\pi^+\pi^-\pi^+e^-$	$CI$	< 3.73	$\times 10^{-4}$	CL=90%
243	$\rho^0 e^+e^-$	$CI$	< 1.0	$\times 10^{-4}$	CL=90%
244	$\pi^+\pi^-\mu^+\mu^-$	$CI$	< 5.5	$\times 10^{-7}$	CL=90%
245	$\rho^0 \mu^+\mu^-$	$CI$	< 2.2	$\times 10^{-5}$	CL=90%
246	$\omega e^+e^-$	$CI$	< 1.8	$\times 10^{-4}$	CL=90%
247	$\omega \mu^+\mu^-$	$CI$	< 8.3	$\times 10^{-4}$	CL=90%
248	$K^-K^+e^+e^-$	$CI$	< 3.15	$\times 10^{-4}$	CL=90%
249	$\phi e^+e^-$	$CI$	< 5.2	$\times 10^{-5}$	CL=90%
250	$K^-K^+\mu^+\mu^-$	$CI$	< 3.3	$\times 10^{-5}$	CL=90%
251	$\phi \mu^+\mu^-$	$CI$	< 3.1	$\times 10^{-5}$	CL=90%
252	$\bar{K}^0 e^+e^-$	$[j]$	< 1.1	$\times 10^{-4}$	CL=90%
253	$\bar{K}^0 \mu^+\mu^-$	$[j]$	< 2.6	$\times 10^{-4}$	CL=90%
254	$K^-\pi^+e^+e^-$	$CI$	< 3.85	$\times 10^{-4}$	CL=90%
255	$\bar{K}^*(892)^0 e^+e^-$	$[j]$	< 4.7	$\times 10^{-5}$	CL=90%
256	$K^-\pi^+\mu^+\mu^-$	$CI$	< 3.59	$\times 10^{-4}$	CL=90%
257	$\bar{K}^*(892)^0 \mu^+\mu^-$	$[j]$	< 2.4	$\times 10^{-5}$	CL=90%
258	$\pi^+\pi^-\pi^0 \mu^+\mu^-$	$CI$	< 8.1	$\times 10^{-4}$	CL=90%
259	$\mu^\pm e^\mp$	$LF$	$[k]$ < 2.6	$\times 10^{-7}$	CL=90%
260	$\pi^0 e^\pm \mu^\mp$	$LF$	$[k]$ < 8.6	$\times 10^{-5}$	CL=90%
261	$\eta e^\pm \mu^\mp$	$LF$	$[k]$ < 1.0	$\times 10^{-4}$	CL=90%
262	$\pi^+\pi^-\pi^\pm \mu^\mp$	$LF$	$[k]$ < 1.5	$\times 10^{-5}$	CL=90%
263	$\rho^0 e^\pm \mu^\mp$	$LF$	$[k]$ < 4.9	$\times 10^{-5}$	CL=90%
264	$\omega e^\pm \mu^\mp$	$LF$	$[k]$ < 1.2	$\times 10^{-4}$	CL=90%
265	$K^-K^+e^\pm \mu^\mp$	$LF$	$[k]$ < 1.8	$\times 10^{-4}$	CL=90%
266	$\phi e^\pm \mu^\mp$	$LF$	$[k]$ < 3.4	$\times 10^{-5}$	CL=90%
267	$\bar{K}^0 e^\pm \mu^\mp$	$LF$	$[k]$ < 1.0	$\times 10^{-4}$	CL=90%
268	$K^-\pi^+e^\pm \mu^\mp$	$LF$	$[k]$ < 5.53	$\times 10^{-4}$	CL=90%
269	$\bar{K}^*(892)^0 e^\pm \mu^\mp$	$LF$	$[k]$ < 8.3	$\times 10^{-5}$	CL=90%
270	$2\pi^-2e^+ + c.c.$	$L$	< 1.12	$\times 10^{-4}$	CL=90%
271	$2\pi^-2\mu^+ + c.c.$	$L$	< 2.9	$\times 10^{-5}$	CL=90%
272	$K^-\pi^-2e^+ + c.c.$	$L$	< 2.06	$\times 10^{-4}$	CL=90%
273	$K^-\pi^-2\mu^+ + c.c.$	$L$	< 3.9	$\times 10^{-4}$	CL=90%
274	$2K^-2e^+ + c.c.$	$L$	< 1.52	$\times 10^{-4}$	CL=90%
275	$2K^-2\mu^+ + c.c.$	$L$	< 9.4	$\times 10^{-5}$	CL=90%
276	$\pi^-\pi^-e^+\mu^+ + c.c.$	$L$	< 7.9	$\times 10^{-5}$	CL=90%
277	$K^-\pi^-e^+\mu^+ + c.c.$	$L$	< 2.18	$\times 10^{-4}$	CL=90%
278	$2K^-e^+\mu^+ + c.c.$	$L$	< 5.7	$\times 10^{-5}$	CL=90%
279	$p e^-$	$L, B$	$[l]$ < 1.0	$\times 10^{-5}$	CL=90%
280	$\bar{p} e^+$	$L, B$	$[n]$ < 1.1	$\times 10^{-5}$	CL=90%

[k] The value is for the sum of the charge states or particle/antiparticle states indicated.

[1] This limit is for either  $D^0$  or  $\bar{D}^0$  to  $p e^-$ .

[n] This limit is for either  $D^0$  or  $\bar{D}^0$  to  $\bar{p}e^+$ .

### CONSTRAINED FIT INFORMATION

An overall fit to 55 branching ratios uses 110 measurements and one constraint to determine 31 parameters. The overall fit has a  $\chi^2 = 108.0$  for 80 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

[illegible]

$\Gamma_{281}$ Unaccounted decay modes	$(37.5 \pm 1.5) \%$	S=1.3
--	---------------------	-------

[a] This value is obtained by subtracting the branching fractions for 2-, 4- and 6-prongs from unity.

[b] This is the sum of our  $K^-2\pi^+\pi^-$ ,  $K^-2\pi^+\pi^-\pi^0$ ,  $\overline{K}^02\pi^+2\pi^-$ ,  $K^+2K^-\pi^+$ ,  $2\pi^+2\pi^-$ ,  $2\pi^+2\pi^-\pi^0$ ,  $K^+K^-\pi^+\pi^-$ , and  $K^+K^-\pi^+\pi^-\pi^0$ , branching fractions.

[c] This is the sum of our  $K^- 3\pi^+ 2\pi^-$  and  $3\pi^+ 3\pi^-$  branching fractions.

[d] The branching fractions for the  $K^- e^+ \nu_e$ ,  $K^*(892)^- e^+ \nu_e$ ,  $\pi^- e^+ \nu_e$ , and  $\rho^- e^+ \nu_e$  modes add up to  $6.19 \pm 0.17$  %.

[e] The branching fraction for this mode may differ from the sum of the submodes that contribute to it, due to interference effects. See the relevant papers.

[f] This is a doubly Cabibbo-suppressed mode.

[g] The two experiments measuring this fraction are in serious disagreement. See the Particle Listings.

[h] Submodes of the  $D^0 \rightarrow K_S^0 \pi^+ \pi^- \pi^0$  mode with a  $K^*$  and/or  $\rho$  were studied by COFFMAN 92B, but with only 140 events. With nothing new for 18 years, we refer to our 2008 edition, Physics Letters **B667** 1 (2008), for those results.

[i] This branching fraction includes all the decay modes of the resonance in the final state.

[j] This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.

# Meson Particle Listings

$D^0$

$x_{173}$	1									
$x_{175}$	1	1								
$x_{176}$	0	0	0							
$x_{177}$	5	4	5	2						
$x_{178}$	0	0	0	0	2					
$x_{179}$	1	0	0	0	3	3				
$x_{181}$	0	0	0	0	3	3	83			
$x_{218}$	1	1	1	0	7	0	0	0		
$x_{222}$	2	1	2	1	12	1	1	1	2	
$x_{281}$	-2	-2	-3	-3	-12	-4	-19	-17	-2	-4
	$x_{167}$	$x_{173}$	$x_{175}$	$x_{176}$	$x_{177}$	$x_{178}$	$x_{179}$	$x_{181}$	$x_{218}$	$x_{222}$

## CONSTRAINED FIT INFORMATION

An overall fit to 3 branching ratios uses 3 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 0.0$  for 0 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-100		
$x_3$	-46	40	
$x_4$	0	0	0
	$x_1$	$x_2$	$x_3$

## $D^0$ BRANCHING RATIOS

Some older now obsolete results have been omitted from these Listings.

### Topological modes

$\Gamma(0\text{-prongs})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
This value is obtained by subtracting the branching fractions for 2-, 4-, and 6-prongs from unity.	
<u>VALUE</u>	<u>DOCUMENT ID</u>
<b><math>0.15 \pm 0.06</math> OUR FIT</b>	

$\Gamma(4\text{-prongs})/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
This is the sum of our $K^- 2\pi^+ \pi^-$ , $K^- 2\pi^+ \pi^- \pi^0$ , $\bar{K}^0 2\pi^+ 2\pi^-$ , $K^+ 2K^- \pi^+$ , $2\pi^+ 2\pi^-$ , $2\pi^+ 2\pi^- \pi^0$ , $K^+ K^- \pi^+ \pi^-$ , and $K^+ K^- \pi^+ \pi^- \pi^0$ branching fractions.	
<u>VALUE</u>	<u>DOCUMENT ID</u>
<b><math>0.145 \pm 0.005</math> OUR FIT</b>	
<b><math>0.145 \pm 0.005</math></b>	PDG 12

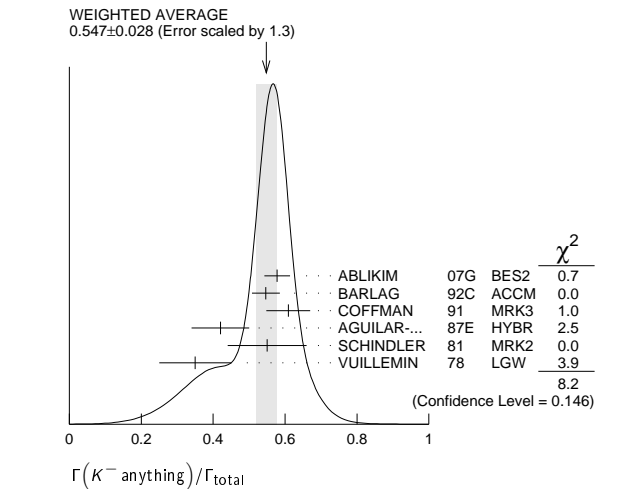
$\Gamma(4\text{-prongs})/\Gamma(2\text{-prongs})$		$\Gamma_3/\Gamma_2$		
<u>VALUE</u>	<u>EVTs</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.207 \pm 0.016</math> OUR FIT</b>				
<b><math>0.207 \pm 0.016 \pm 0.004</math></b>	226	ONENGUT	05	CHRS $\nu_\mu$ emulsion, $\overline{E}_\nu \approx 27$ GeV

(6-prongs)/ $\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
This is the sum of our $K^- 3\pi^+ 2\pi^-$ and $3\pi^+ 3\pi^-$ branching fractions.				
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTs</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>6.4 <math>\pm</math> 1.3 OUR FIT</b>				
<b>6.4 <math>\pm</math> 1.3</b>		PDG	12	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
12 $\begin{smallmatrix} +13 \\ -9 \end{smallmatrix} \pm 2$	3	ONENGUT	05	CHRS $\nu_\mu$ emulsion, $\bar{E}_\nu \approx 27$ GeV

### Inclusive modes

$\Gamma(e^+ \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$			
The branching fractions for the $K^- e^+ \nu_e$ , $K^*(892)^- e^+ \nu_e$ , $\pi^- e^+ \nu_e$ , and $\rho^- e^+ \nu_e$ modes add up to $6.20 \pm 0.17$ %.				
VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>6.49 \pm 0.11</math> OUR AVERAGE</b>				
$6.46 \pm 0.09 \pm 0.11$	$6584 \pm 96$	<sup>1</sup> ASNER	10	CLEO $e^+ e^-$ at 3774 MeV
$6.3 \pm 0.7 \pm 0.4$	$290 \pm 32$	ABLIKIM	07G	BES2 $e^+ e^- \approx \psi(3770)$
$6.46 \pm 0.17 \pm 0.13$	$2246 \pm 57$	ADAM	06A	CLEO See ASNER 10
$6.9 \pm 0.3 \pm 0.5$	1670	ALBRECHT	96c	ARG $e^+ e^- \approx 10$ GeV
$6.64 \pm 0.18 \pm 0.29$	4609	KUBOTA	96B	CLE2 $e^+ e^- \approx \Upsilon(4S)$
<sup>1</sup> Using the $D^+$ and $D^0$ lifetimes, ASNER 10 finds that the ratio of the $D^+$ and $D^0$ semileptonic widths is $0.985 \pm 0.015 \pm 0.024$ .				
$\Gamma(\mu^+ \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$			
VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>6.7 \pm 0.6</math> OUR FIT</b>				
<b><math>6.4 \pm 0.8</math> OUR AVERAGE</b>				
$6.8 \pm 1.5 \pm 0.8$	$79 \pm 10$	<sup>1</sup> ABLIKIM	08L	BES2 $e^+ e^- \approx \psi(3772)$
$6.5 \pm 1.2 \pm 0.3$	36	KAYIS-TOPAK.05	CHRS	$\nu_\mu$ emulsion
$6.0 \pm 0.7 \pm 1.2$	310	ALBRECHT	96c	ARG $e^+ e^- \approx 10$ GeV
<sup>1</sup> ABLIKIM 08L finds the ratio of $D^+ \rightarrow \mu^+ X$ and $D^0 \rightarrow \mu^+ X$ branching fractions to be $2.59 \pm 0.70 \pm 0.25$ , in accord with the ratio of $D^+$ and $D^0$ lifetimes, $2.54 \pm 0.02$ .				

$\Gamma(K^- \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.547 \pm 0.028</math> OUR AVERAGE</b>	Error	includes scale factor of 1.3. See the ideogram below.		
$0.578 \pm 0.016 \pm 0.032$	$2098 \pm 59$	ABLIKIM	07G	BES2 $e^+e^- \approx \psi(3770)$
$0.546^{+0.039}_{-0.038}$		<sup>1</sup> BARLAG	92c	ACCM $\pi^-$ Cu 230 GeV
$0.609 \pm 0.032 \pm 0.052$		COFFMAN	91	MRK3 $e^+e^-$ 3.77 GeV
$0.42 \pm 0.08$		AGUILAR-...	87E	HYBR $\pi p, pp$ 360, 400 GeV
$0.55 \pm 0.11$	121	SCHINDLER	81	MRK2 $e^+e^-$ 3.771 GeV
$0.35 \pm 0.10$	19	VUILLEMIN	78	LGW $e^+e^-$ 3.772 GeV
<sup>1</sup> BARLAG 92c computes the branching fraction using topological normalization.				



$[\Gamma(K^0 \text{ anything}) + \Gamma(K^0 \text{ anything})]/\Gamma_{\text{total}}$			$\Gamma_8/\Gamma$	
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.47 \pm 0.04</math> OUR AVERAGE</b>				
$0.476 \pm 0.048 \pm 0.030$	$250 \pm 25$	ABLIKIM	06u	BES2 $e^+e^-$ at 3773 MeV
$0.455 \pm 0.050 \pm 0.032$		COFFMAN	91	MRK3 $e^+e^-$ 3.77 GeV

$\Gamma(K^+ \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.034 \pm 0.004</math> OUR AVERAGE</b>				
$0.035 \pm 0.007 \pm 0.003$	$119 \pm 23$	ABLIKIM	07G BES2	$e^+e^- \approx \psi(3770)$
$0.034^{+0.007}_{-0.005}$		<sup>1</sup> BARLAG	92c ACCM	$\pi^-$ Cu 230 GeV
$0.028 \pm 0.009 \pm 0.004$		COFFMAN	91 MRK3	$e^+e^-$ 3.77 GeV
$0.03^{+0.05}_{-0.02}$		AGUILAR-...	87E HYBR	$\pi p, pp$ 360, 400 GeV
$0.08 \pm 0.03$	25	SCHINDLER	81 MRK2	$e^+e^-$ 3.771 GeV
<sup>1</sup> BARLAG 92c computes the branching fraction using topological normalization.				

$\Gamma(K^*(892)^- \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
<u>VALUE</u>	<u>EVTs</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.153 \pm 0.083 \pm 0.019</math></b>	$28 \pm 15$	ABLIKIM	06u BES2	$e^+ e^-$ at 3773 MeV

$\Gamma(\bar{K}^*(892)^0 \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.087 \pm 0.040 \pm 0.012</math></b>	$96 \pm 44$	ABLIKIM	05P BES	$e^+e^- \approx 3773 \text{ MeV}$

$\Gamma(K^*(892)+\text{anything})/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.036</b>	90	ABLIKIM	06u	BES2	$e^+e^-$ at 3773 MeV

$\Gamma(K^*(892)^0 \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.028 \pm 0.012 \pm 0.004</math></b>	$31 \pm 12$	ABLIKIM	05P BES	$e^+e^- \approx 3773 \text{ MeV}$

$\Gamma(\eta \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
This ratio includes $\eta$ particles from $\eta'$ decays.				
VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>9.5 \pm 0.4 \pm 0.8</math></b>	$4463 \pm 197$	HUANG	06B CLEO	$e^+e^-$ at $\psi(3770)$

$\Gamma(\eta' \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{15}/\Gamma$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTs</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>2.48 \pm 0.17 \pm 0.21</math></b>	$299 \pm 21$	HUANG	06B CLEO	$e^+ e^-$ at $\psi(3770)$

$\Gamma(\phi \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{16}/\Gamma$
VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>1.05 \pm 0.08 \pm 0.07</math></b>	$368 \pm 24$	HUANG	06B	CLEO $e^+ e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.71^{+0.76}_{-0.71} \pm 0.17$	9	BAI	00c	BES $e^+ e^- \rightarrow D \bar{D}^*, D^* \bar{D}^*$

# Meson Particle Listings

 $D^0$ 

Semileptonic modes					$\Gamma_{18}/\Gamma$	
$\Gamma(K^- e^+ \nu_e)/\Gamma_{\text{total}}$						
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>3.538 \pm 0.033</math> OUR FIT</b>	Error	includes scale factor of 1.3.				
<b><math>3.503 \pm 0.029</math> OUR AVERAGE</b>						
$3.505 \pm 0.014 \pm 0.033$	71k	<sup>1</sup> ABLIKIM	15x BES3	$2.92 \text{ fb}^{-1}$ , 3.773 GeV		
$3.50 \pm 0.03 \pm 0.04$	14.1k	<sup>1</sup> BESSON	09 CLEO	$e^+ e^-$ at $\psi(3770)$		
$3.45 \pm 0.10 \pm 0.19$	1.3k	<sup>2</sup> WIDHALM	06 BELL	$e^+ e^- \approx \Upsilon(4S)$		
$3.82 \pm 0.40 \pm 0.27$	104	ABLIKIM	04c BES	$e^+ e^-$ , 3.773 GeV		
$3.4 \pm 0.5 \pm 0.4$	55	ADLER	89 MRK3	$e^+ e^-$ 3.77 GeV		
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
$3.56 \pm 0.03 \pm 0.09$		<sup>3</sup> DOBBS	08 CLEO	See BESSON 09		
$3.44 \pm 0.10 \pm 0.10$	1.3k	COAN	05 CLEO	See DOBBS 08		

<sup>1</sup> See the form-factor parameters near the end of this  $D^0$  Listing.

<sup>2</sup> The  $\pi^- e^+ \nu_e$  and  $K^- e^+ \nu_e$  results of WIDHALM 06 give  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_{\pi^+}(0)}{f_K^+(0)}|^2 = 0.042 \pm 0.003 \pm 0.003$ .

<sup>3</sup> DOBBS 08 establishes  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_{\pi^+}(0)}{f_K^+(0)}| = 0.188 \pm 0.008 \pm 0.002$  from the  $D^+$  and  $D^0$  decays to  $\overline{K} e^+ \nu_e$  and  $\pi e^+ \nu_e$ .

<sup>1</sup> See the form-factor parameters near the end of this  $D^0$  Listing.

<sup>2</sup> The  $\pi^- e^+ \nu_e$  and  $K^- e^+ \nu_e$  results of WIDHALM 06 give  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.042 \pm 0.003 \pm 0.003$ .

<sup>3</sup> DOBBS 08 establishes  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}| = 0.188 \pm 0.008 \pm 0.002$  from the  $D^+$  and  $D^0$  decays to  $\bar{K} e^+ \nu_e$  and  $\pi e^+ \nu_e$ .

$\Gamma(K^- e^+ \nu_e)/\Gamma(K^- \pi^+)$					$\Gamma_{18}/\Gamma_{31}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.900±0.012 OUR FIT</b>	Error	includes scale factor of 1.3.			
<b>0.930±0.013 OUR AVERAGE</b>					
0.927±0.007±0.012	76k±323	<sup>1</sup> AUBERT	07Bg BABR	$e^+ e^- \approx \gamma(4S)$	
0.978±0.027±0.044	2510	<sup>2</sup> BEAN	93c CLE2	$e^+ e^- \approx \gamma(4S)$	
0.90±0.06±0.06	584	<sup>3</sup> CRAWFORD	91B CLEO	$e^+ e^- \approx 10.5 \text{ GeV}$	
0.91±0.07±0.11	250	<sup>4</sup> ANJOS	89F E691	Photoproduction	

<sup>1</sup> The event samples in this AUBERT 07Bg result include radiative photons. The  $D^0 \rightarrow K^- e^+ \nu_e$  form factor at  $q^2 = 0$  is  $f_+(0) = 0.727 \pm 0.007 \pm 0.005 \pm 0.007$ .

<sup>2</sup> BEAN 93c uses  $K^- \mu^+ \nu_\mu$  as well as  $K^- e^+ \nu_e$  events and makes a small phase-space adjustment to the number of the  $\mu^+$  events to use them as  $e^+$  events. A pole mass of  $2.00 \pm 0.12 \pm 0.18 \text{ GeV}/c^2$  is obtained from the  $q^2$  dependence of the decay rate.

<sup>3</sup> CRAWFORD 91B uses  $K^- e^+ \nu_e$  and  $K^- \mu^+ \nu_\mu$  candidates to measure a pole mass of  $2.1^{+0.4+0.3}_{-0.2-0.2} \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

<sup>4</sup> ANJOS 89F measures a pole mass of  $2.1^{+0.4}_{-0.2} \pm 0.2 \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(K^- \mu^+ \nu_\mu)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.33 \pm 0.13</math> OUR FIT</b>					
<b><math>3.45 \pm 0.10 \pm 0.21</math></b>	1249 $\pm$ 43	WIDHALM	06 BELL	$e^+ e^- \approx \mathcal{T}(4S)$	

$\Gamma(K^- \mu^+ \nu_\mu)/\Gamma(K^- \pi^+)$					$\Gamma_{19}/\Gamma_{31}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.848 \pm 0.033</math> OUR FIT</b>					
<b><math>0.84 \pm 0.04</math> OUR AVERAGE</b>					
$0.852 \pm 0.034 \pm 0.028$	1897	<sup>1</sup> FRABETTI	95G E687	$\gamma \text{ Be } \bar{E}_\gamma = 220 \text{ GeV}$	
$0.82 \pm 0.13 \pm 0.13$	338	<sup>2</sup> FRABETTI	93i E687	$\gamma \text{ Be } \bar{E}_\gamma = 221 \text{ GeV}$	
$0.79 \pm 0.08 \pm 0.09$	231	<sup>3</sup> CRAWFORD	91B CLEO	$e^+ e^- \approx 10.5 \text{ GeV}$	

<sup>1</sup> FRABETTI 95G extracts the ratio of form factors  $f_-(0)/f_+(0) = -1.3^{+3.6}_{-0.8-0.06} \pm 0.6$ , and measures a pole mass of  $1.87^{+0.11+0.07}_{-0.08-0.06} \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

<sup>2</sup> FRABETTI 93i measures a pole mass of  $2.1^{+0.7+0.7}_{-0.3-0.3} \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

<sup>3</sup> CRAWFORD 91B measures a pole mass of  $2.00 \pm 0.12 \pm 0.18 \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(K^- \mu^+ \nu_\mu)/\Gamma(\mu^+ \text{ anything})$					$\Gamma_{19}/\Gamma_6$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.50 \pm 0.05</math> OUR FIT</b>					
<b><math>0.472 \pm 0.051 \pm 0.040</math></b>	232	KODAMA	94 E653	$\pi^-$ emulsion 600 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.32 \pm 0.05 \pm 0.05$	124	KODAMA	91 EMUL	$pA$ 800 GeV	

$\Gamma(K^- \pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.016^{+0.013}_{-0.005} \pm 0.002</math></b>	4	<sup>1</sup> BAI	91 MRK3	$e^+ e^- \approx 3.77 \text{ GeV}$	

<sup>1</sup> BAI 91 finds that a fraction  $0.79^{+0.15+0.09}_{-0.17-0.03}$  of combined  $D^+$  and  $D^0$  decays to  $\bar{K} \pi e^+ \nu_e$  (24 events) are  $\bar{K}^*(892) e^+ \nu_e$ . BAI 91 uses 56  $K^- e^+ \nu_e$  events to measure a pole mass of  $1.8 \pm 0.3 \pm 0.2 \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(\bar{K}^0 \pi^- e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.7^{+0.9}_{-0.7}</math> OUR AVERAGE</b>					
$2.61 \pm 1.04 \pm 0.28$	9 $\pm$ 3	ABLIKIM	06o BES2	$e^+ e^-$ at 3773 MeV	
$2.8^{+1.7}_{-0.8} \pm 0.3$	6	<sup>1</sup> BAI	91 MRK3	$e^+ e^- \approx 3.77 \text{ GeV}$	

<sup>1</sup> BAI 91 finds that a fraction  $0.79^{+0.15+0.09}_{-0.17-0.03}$  of combined  $D^+$  and  $D^0$  decays to  $\bar{K} \pi e^+ \nu_e$  (24 events) are  $\bar{K}^*(892) e^+ \nu_e$ .

$\Gamma(K^*(892)^- e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$
Both decay modes of the $K^*(892)^-$ are included.					
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.16±0.16 OUR FIT</b>					
<b>2.16±0.15±0.08</b>	219 ± 16	<sup>1</sup> COAN	05 CLEO	$e^+ e^-$ at $\psi(3770)$	
<sup>1</sup> COAN 05 uses both $K^-\pi^0$ and $K_S^0\pi^-$ events.					

$\Gamma(K^*(892)^- e^+ \nu_e)/\Gamma(K_S^0 \pi^+ \pi^-)$					$\Gamma_{20}/\Gamma_{35}$
Unseen decay modes of the $K^*(892)^-$ are included.					
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b><math>0.76 \pm 0.07</math> OUR FIT</b>					
<b><math>0.76 \pm 0.12 \pm 0.06</math></b>	152	<sup>1</sup> BEAN	93c	$e^+ e^- \approx \gamma(4S)$	
<sup>1</sup> BEAN 93c uses $K^{*-} \mu^+ \nu_\mu$ as well as $K^{*-} e^+ \nu_e$ events and makes a small phase-space adjustment to the number of the $\mu^+$ events to use them as $e^+$ events.					

$\Gamma(K^*(892)^-\mu^+\nu_\mu)/\Gamma(K_S^0\pi^+\pi^-)$					$\Gamma_{21}/\Gamma_{35}$
Unseen decay modes of the $K^*(892)^-$ are included.					
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.674 \pm 0.068 \pm 0.026</math></b>	$175 \pm 17$	<sup>1</sup> LINK	05B FOCS	$\gamma A, \bar{E}_\gamma \approx 180 \text{ GeV}$	
<sup>1</sup> LINK 05B finds that in $D^0 \rightarrow \bar{K}^0\pi^-\mu^+\nu_\mu$ the $\bar{K}^0\pi^-$ system is 6% in S-wave.					

$\Gamma(K^- \pi^+ \pi^- e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.8^{+1.4}_{-1.1} \pm 0.3</math></b>	8	ARTUSO	07A CLEO	$e^+ e^-$ at $\mathcal{T}(3770)$	

$\Gamma(K_1(1270)^- e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>7.6^{+4.1}_{-3.0} \pm 0.9</math></b>	8	<sup>1</sup> ARTUSO	07A CLEO	$e^+ e^-$ at $\mathcal{T}(3770)$	
<sup>1</sup> This ARTUSO 07A result is corrected for all decay modes of the $K_1(1270)^-$ .					

$\Gamma(K^- \pi^+ \pi^- \mu^+ \nu_\mu)/\Gamma(K^- \mu^+ \nu_\mu)$					$\Gamma_{26}/\Gamma_{19}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;0.037</math></b>	90	KODAMA	93B E653	$\pi^-$ emulsion 600 GeV	

$\Gamma((\bar{K}^*(892) \pi)^- \mu^+ \nu_\mu)/\Gamma(K^- \mu^+ \nu_\mu)$					$\Gamma_{27}/\Gamma_{19}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;0.043</math></b>	90	<sup>1</sup> KODAMA	93B E653	$\pi^-$ emulsion 600 GeV	
<sup>1</sup> KODAMA 93B searched in $K^- \pi^+ \pi^- \mu^+ \nu_\mu$ , but the limit includes other $(\bar{K}^*(892) \pi)^-$ charge states.					

$\Gamma(\pi^- e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.291 \pm 0.004</math> OUR FIT</b>	Error	includes scale factor of 1.1.			
<b><math>0.293 \pm 0.004</math> OUR AVERAGE</b>					
$0.295 \pm 0.004 \pm 0.003$	6.3k	<sup>1</sup> ABLIKIM	15x BES3	$2.92 \text{ fb}^{-1}$ , 3.773 GeV	
$0.288 \pm 0.008 \pm 0.003$	1.3k	<sup>1</sup> BESSON	09 CLEO	$e^+ e^-$ at $\psi(3770)$	
$0.279 \pm 0.027 \pm 0.016$	126	<sup>2</sup> WIDHALM	06 BELL	$e^+ e^- \approx \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.299 \pm 0.011 \pm 0.009$		<sup>3</sup> DOBBS	08 CLEO	See BESSON 09	
$0.262 \pm 0.025 \pm 0.008$	117	COAN	05 CLEO	See DOBBS 08	

<sup>1</sup> See the form-factor parameters near the end of this  $D^0$  Listing.

<sup>2</sup> The  $\pi^- e^+ \nu_e$  and  $K^- e^+ \nu_e$  results of WIDHALM 06 give  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.042 \pm 0.003 \pm 0.003$ .

<sup>3</sup> DOBBS 08 establishes  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}| = 0.188 \pm 0.008 \pm 0.002$  from the  $D^+$  and  $D^0$  decays to  $\bar{K} e^+ \nu_e$  and  $\pi e^+ \nu_e$ .

$\Gamma(\pi^- e^+ \nu_e)/\Gamma(K^- e^+ \nu_e)$					$\Gamma_{28}/\Gamma_{18}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
$0.0823 \pm 0.0014$ OUR FIT Error includes scale factor of 1.1.					
$0.085 \pm 0.007$ OUR AVERAGE					
$0.082 \pm 0.006 \pm 0.005$		<sup>1</sup> HUANG	05 CLEO	$e^+ e^- \approx \Upsilon(4S)$	
$0.101 \pm 0.020 \pm 0.003$	91	<sup>2</sup> FRABETTI	96B E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV	
$0.103 \pm 0.039 \pm 0.013$	87	<sup>3</sup> BUTLER	95 CLE2	$< 0.156$ (90% CL)	

<sup>1</sup> HUANG 05 uses both  $e$  and  $\mu$  events, and makes a small correction to the  $\mu$  events to make them effectively  $e$  events. This result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.038^{+0.006+0.005}_{-0.007-0.003}$ .

<sup>2</sup> FRABETTI 96B uses both  $e$  and  $\mu$  events, and makes a small correction to the  $\mu$  events to make them effectively  $e$  events. This result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.050 \pm 0.011 \pm 0.002$ .

<sup>3</sup> BUTLER 95 has 87  $\pm$  33  $\pi^- e^+ \nu_e$  events. The result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.052 \pm 0.020 \pm 0.007$ .

# Meson Particle Listings

$D^0$

$\Gamma(\pi^- e^+ \nu_e)/\Gamma(K^- \pi^+)$					$\Gamma_{28}/\Gamma_{31}$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>7.41±0.13 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>7.02±0.17±0.23</b>	375k	<sup>1</sup> LEES	15F BABR	347 fb <sup>-1</sup> , 10.58 GeV	
<sup>1</sup> See the form-factor parameters near the end of the $D^0$ Listing.					

$\Gamma(\pi^- \mu^+ \nu_\mu)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.238±0.024 OUR FIT</b>					
<b>0.231±0.026±0.019</b>	106 ± 13	WIDHALM	06 BELL	$e^+ e^- \approx \mathcal{T}(4S)$	

$\Gamma(\pi^- \mu^+ \nu_\mu)/\Gamma(K^- \mu^+ \nu_\mu)$					$\Gamma_{29}/\Gamma_{19}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.071±0.007 OUR FIT</b>					
<b>0.074±0.008±0.007</b>	288 ± 29	<sup>1</sup> LINK	05 FOCS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV	
<sup>1</sup> LINK 05 finds the form-factor ratio $ f_0^\pi(0)/f_0^K(0) $ to be $0.85 \pm 0.04 \pm 0.04 \pm 0.01$ .					

$\Gamma(\rho^- e^+ \nu_e)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.77±0.12±0.10</b>	305 ± 21	<sup>1,2</sup> DOBBS	13 CLEO	$e^+ e^-$ at $\psi(3770)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.94±0.39±0.13	31 ± 6	COAN	05 CLEO	See DOBBS 13	
<sup>1</sup> DOBBS 13 finds $\Gamma(D^0 \rightarrow \rho^- e^+ \nu_e) / 2 \Gamma(D^+ \rightarrow \rho^0 e^+ \nu_e) = 1.03 \pm 0.09^{+0.08}_{-0.02}$ ; isospin invariance predicts the ratio is 1.0.					
<sup>2</sup> See the $D^+$ Listings for $D \rightarrow \rho e^+ \nu_e$ form factors.					

## Hadronic modes with a single $\overline{K}$

$\Gamma(K^- \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3.93 ±0.04 OUR FIT</b>	Error includes scale factor of 1.2.				
<b>3.93 ±0.05 OUR AVERAGE</b>	Error includes scale factor of 1.1.				
3.934±0.021±0.061		BONVICINI	14 CLEO	All CLEO-c runs	
4.007±0.037±0.072	33.8k	AUBERT	08L BABR	$e^+ e^-$ at $\mathcal{T}(4S)$	
3.82 ±0.07 ±0.12		<sup>1</sup> ARTUSO	98 CLE2	CLEO average	
3.90 ±0.09 ±0.12	5.4k	<sup>2</sup> BARATE	97c ALEP	From Z decays	
3.41 ±0.12 ±0.28	1.2k	<sup>2</sup> ALBRECHT	94F ARG	$e^+ e^- \approx \mathcal{T}(4S)$	
3.62 ±0.34 ±0.44		<sup>2</sup> DECAMP	91J ALEP	From Z decays	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.891±0.035±0.069		<sup>3</sup> DOBBS	07 CLEO	See BONVICINI 14	
3.91 ±0.08 ±0.09	10.3k	<sup>3</sup> HE	05 CLEO	See DOBBS 07	
3.81 ±0.15 ±0.16	1.2k	<sup>4</sup> ARTUSO	98 CLE2	$e^+ e^-$ at $\mathcal{T}(4S)$	
3.69 ±0.11 ±0.16		<sup>5</sup> COAN	98 CLE2	See ARTUSO 98	
4.5 ±0.6 ±0.4		<sup>6</sup> ALBRECHT	94 ARG	$e^+ e^- \approx \mathcal{T}(4S)$	
3.95 ±0.08 ±0.17	4.2k	<sup>2,7</sup> AKERIB	93 CLE2	See ARTUSO 98	
4.5 ±0.8 ±0.5	56	<sup>2</sup> ABACHI	88 HRS	$e^+ e^-$ 29 GeV	
4.2 ±0.4 ±0.4	0.9k	ADLER	88c MRK3	$e^+ e^-$ 3.77 GeV	
4.1 ±0.6	0.3k	<sup>8</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV	
4.3 ±1.0	130	<sup>9</sup> PERUZZI	77 LGW	$e^+ e^-$ 3.77 GeV	

- <sup>1</sup> This combines the CLEO results of ARTUSO 98, COAN 98, and AKERIB 93.
- <sup>2</sup> ABACHI 88, DECAMP 91J, AKERIB 93, ALBRECHT 94F, and BARATE 97c use  $D^*(2010)^+ \rightarrow D^0 \pi^+$  decays. The  $\pi^+$  is both slow and of low  $p_T$  with respect to the event thrust axis or nearest jet ( $\approx D^{*+}$  direction). The excess number of such  $\pi^+$ 's over background gives the number of  $D^*(2010)^+ \rightarrow D^0 \pi^+$  events, and the fraction with  $D^0 \rightarrow K^- \pi^+$  gives the  $D^0 \rightarrow K^- \pi^+$  branching fraction.
- <sup>3</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.
- <sup>4</sup> ARTUSO 98, following ALBRECHT 94, uses  $D^0$  mesons from  $\overline{B}^0 \rightarrow D^*(2010)^+ X \ell^- \overline{\nu}_\ell$  decays. Our average uses the CLEO average of this value with the values of COAN 98 and AKERIB 93.
- <sup>5</sup> COAN 98 assumes that  $\Gamma(B \rightarrow \overline{D} X \ell^+ \nu)/\Gamma(B \rightarrow X \ell^+ \nu) = 1.0 - 3|V_{ub}/V_{cb}|^2 - 0.010 \pm 0.005$ , the last term accounting for  $\overline{B} \rightarrow D_s^+ K X \ell^- \overline{\nu}$ . COAN 98 is included in the CLEO average in ARTUSO 98.
- <sup>6</sup> ALBRECHT 94 uses  $D^0$  mesons from  $\overline{B}^0 \rightarrow D^{*+} \ell^- \overline{\nu}_\ell$  decays. This is a different set of events than used by ALBRECHT 94F.
- <sup>7</sup> This AKERIB 93 value includes radiative corrections; without them, the value is  $0.0391 \pm 0.0008 \pm 0.0017$ . AKERIB 93 is included in the CLEO average in ARTUSO 98.
- <sup>8</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.24 \pm 0.02$  nb. We use the MARK-3 (ADLER 88c) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.
- <sup>9</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.25 \pm 0.05$  nb. We use the MARK-3 (ADLER 88c) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

$\Gamma(K^+ \pi^-)/\Gamma(K^- \pi^+)$				$\Gamma_{32}/\Gamma_{31}$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>3.56 ±0.06 OUR AVERAGE</b>				
3.53 ±0.13	<sup>1</sup> KO	14 BELL	$e^+ e^- \rightarrow \mathcal{T}(nS)$	
3.568±0.066	<sup>2</sup> AAIJ	13CE LHCB	$p\overline{p}$ at 7, 8 TeV	
3.51 ±0.35	<sup>3</sup> AALTONEN	13AE CDF	$p\overline{p}$ at 1.96 TeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.52 ±0.15	<sup>4</sup> AAIJ	13N LHCB	Repl. by AAIJ 13Ce	

- <sup>1</sup> Based on 976 fb<sup>-1</sup> of data collected at  $Y(nS)$  resonances. Assumes no  $CP$  violation.
- <sup>2</sup> Based on 3 fb<sup>-1</sup> of data collected at  $\sqrt{s} = 7, 8$  TeV. Assumes no  $CP$  violation.
- <sup>3</sup> Based on 9.6 fb<sup>-1</sup> of data collected at the Tevatron. Assumes no  $CP$  violation.
- <sup>4</sup> Based on 1 fb<sup>-1</sup> of data collected at  $\sqrt{s} = 7$  TeV in 2011. Assumes no  $CP$  violation.

$\Gamma(K_S^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.240±0.017±0.056	614	HE	08 CLEO	See MENDEZ 10	

$\Gamma(K_S^0 \pi^0)/\Gamma(K^- \pi^+)$					$\Gamma_{33}/\Gamma_{31}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.68±0.12±0.11	119	ANJOS	92B E691	$\gamma$ Be 80–240 GeV	

$\Gamma(K_S^0 \pi^0)/[\Gamma(K^- \pi^+) + \Gamma(K^+ \pi^-)]$					$\Gamma_{33}/(\Gamma_{31} + \Gamma_{222})$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>30.5±0.9 OUR FIT</b>					
<b>30.4±0.3±0.9</b>	20k	MENDEZ	10 CLEO	$e^+ e^-$ at 3774 MeV	

$\Gamma(K_S^0 \pi^0)/\Gamma(K_S^0 \pi^+ \pi^-)$					$\Gamma_{33}/\Gamma_{35}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.421 ±0.029 OUR FIT</b>					
<b>0.44 ±0.02 ±0.05</b>	1942 ± 64	PROCARIO	93B CLE2	$e^+ e^-$ 10.36–10.7 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.34 ±0.04 ±0.02	92	<sup>1</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV	
0.36 ±0.04 ±0.08	104	KINOSHITA	91 CLEO	$e^+ e^- \sim 10.7$ GeV	

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K_L^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{34}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.998±0.049±0.048</b>	1116	<sup>1</sup> HE	08 CLEO	$e^+ e^-$ at $\psi(3770)$	
<sup>1</sup> The difference of HE 08 $D^0 \rightarrow K_S^0 \pi^0$ and $K_L^0 \pi^0$ branching fractions over the sum is $0.108 \pm 0.025 \pm 0.024$ . This is consistent with U-spin symmetry and the Cabibbo angle.					

$\Gamma(K_S^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{35}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.52±0.20±0.25	284 ± 22	<sup>1</sup> ALBRECHT	94F ARG	$e^+ e^- \approx \mathcal{T}(4S)$	
3.2 ±0.3 ±0.5		ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV	
2.6 ±0.8	32 ± 8	<sup>2</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV	
4.0 ±1.2	28	<sup>3</sup> PERUZZI	77 LGW	$e^+ e^-$ 3.77 GeV	
<sup>1</sup> See the footnote on the ALBRECHT 94F measurement of $\Gamma(K^- \pi^+)/\Gamma_{\text{total}}$ for the method used.					
<sup>2</sup> SCHINDLER 81 (MARK-2) measures $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$ branching fraction to be $0.30 \pm 0.08$ nb. We use the MARK-3 (ADLER 88c) value of $\sigma = 5.8 \pm 0.5 \pm 0.6$ nb.					
<sup>3</sup> PERUZZI 77 (MARK-1) measures $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$ branching fraction to be $0.46 \pm 0.12$ nb. We use the MARK-3 (ADLER 88c) value of $\sigma = 5.8 \pm 0.5 \pm 0.6$ nb.					

$\Gamma(K_S^0 \pi^+ \pi^-)/\Gamma(K^- \pi^+)$					$\Gamma_{35}/\Gamma_{31}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.73±0.05 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>0.81±0.05±0.08</b>	856 ± 35	FRABETTI	94J E687	$\gamma$ Be $\overline{E}_\gamma=220$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.85±0.40	35	AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$	
1.4 ±0.5	116	PICCOLO	77 MRK1	$e^+ e^-$ 4.03, 4.41 GeV	

$\Gamma(K_S^0 \rho^0)/\Gamma(K_S^0 \pi^+ \pi^-)$					$\Gamma_{36}/\Gamma_{35}$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.224 ±0.017 −0.023 OUR AVERAGE</b>	Error includes scale factor of 1.7.				
0.210±0.016	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts		
0.264±0.009 ±0.010 −0.026	MURAMATSU	02 CLE2	Dalitz fit, 5299 evts		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.267±0.011 ±0.009 −0.028	ASNER	04A CLEO	See MURAMATSU 02		
0.350±0.028±0.067	FRABETTI	94G E687	Dalitz fit, 597 evts		
0.227±0.032±0.009	ALBRECHT	93D ARG	Dalitz fit, 440 evts		
0.215±0.051±0.037	ANJOS	93 E691	$\gamma$ Be 90–260 GeV		
0.20 ±0.06 ±0.03	FRABETTI	92B E687	$\gamma$ Be, $\overline{E}_\gamma=221$ GeV		
0.12 ±0.01 ±0.07	ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV		

<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

$\Gamma(K_S^0 \omega, \omega \rightarrow \pi^+ \pi^-)/\Gamma(K_S^0 \pi^+ \pi^-)$					$\Gamma_{37}/\Gamma_{35}$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.0073±0.0020 OUR AVERAGE</b>					
0.009 ±0.010	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts		
0.0072±0.0018 ±0.0010 −0.0009	MURAMATSU	02 CLE2	Dalitz fit, 5299 evts		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.0081±0.0019 ±0.0018 −0.0010	ASNER	04A CLEO	See MURAMATSU 02		

<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

See key on page 601

## Meson Particle Listings

 $D^0$  $\Gamma(K_S^0(\pi^+\pi^-)_{S\text{-wave}})/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{38}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis. The  $(\pi^+\pi^-)_{S\text{-wave}}$  includes what in isobar models are the  $f_0(980)$  and  $f_0(1370)$ ; see the following two data blocks.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.119±0.026</b>	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts

<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K_S^0 f_0(980), f_0(980) \rightarrow \pi^+\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{39}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.043±0.005<sup>+0.012</sup><sub>-0.006</sub></b>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.042±0.005 <sup>+0.011</sup> <sub>-0.005</sub>	ASNER	04A CLEO	See MURAMATSU 02
0.068±0.016±0.018	FRABETTI	94G E687	Dalitz fit, 597 evts
0.046±0.018±0.006	ALBRECHT	93D ARG	Dalitz fit, 440 evts

 $\Gamma(K_S^0 f_0(1370), f_0(1370) \rightarrow \pi^+\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{40}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.099±0.011<sup>+0.028</sup><sub>-0.044</sub></b>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.098±0.014 <sup>+0.026</sup> <sub>-0.036</sub>	ASNER	04A CLEO	See MURAMATSU 02
0.077±0.022±0.031	FRABETTI	94G E687	Dalitz fit, 597 evts
0.082±0.028±0.013	ALBRECHT	93D ARG	Dalitz fit, 440 evts

 $\Gamma(K_S^0 f_2(1270), f_2(1270) \rightarrow \pi^+\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{41}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0032<sup>+0.0035</sup><sub>-0.0022</sub></b> OUR AVERAGE			

0.006 ±0.007	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts
0.0027±0.0015 <sup>+0.0037</sup> <sub>-0.0017</sub>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.0036±0.0022 <sup>+0.0032</sup> <sub>-0.0019</sub>	ASNER	04A CLEO	See MURAMATSU 02
0.037 ±0.014 ±0.017	FRABETTI	94G E687	Dalitz fit, 597 evts
0.050 ±0.021 ±0.008	ALBRECHT	93D ARG	Dalitz fit, 440 evts

<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K^*(892)^-\pi^+, K^*(892)^-\rightarrow K_S^0\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{42}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.588<sup>+0.034</sup><sub>-0.050</sub></b> OUR AVERAGE	Error includes scale factor of 2.0.		

0.557±0.028	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts
0.657±0.013 <sup>+0.018</sup> <sub>-0.040</sub>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.663±0.013 <sup>+0.024</sup> <sub>-0.043</sub>	ASNER	04A CLEO	See MURAMATSU 02
0.625±0.036±0.026	FRABETTI	94G E687	Dalitz fit, 597 evts
0.718±0.042±0.030	ALBRECHT	93D ARG	Dalitz fit, 440 evts
0.480±0.097	ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.56 ±0.04 ±0.05	ADLER	87 MRK3	$e^+e^-$ 3.77 GeV

<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K_0^*(1430)^-\pi^+, K_0^*(1430)^-\rightarrow K_S^0\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{43}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.095<sup>+0.014</sup><sub>-0.010</sub></b> OUR AVERAGE			

0.102±0.015	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts
0.073±0.007 <sup>+0.031</sup> <sub>-0.011</sub>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.072±0.007 <sup>+0.014</sup> <sub>-0.013</sub>	ASNER	04A CLEO	See MURAMATSU 02
0.109±0.027±0.029	FRABETTI	94G E687	Dalitz fit, 597 evts
0.129±0.034±0.021	ALBRECHT	93D ARG	Dalitz fit, 440 evts

<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K_2^*(1430)^-\pi^+, K_2^*(1430)^-\rightarrow K_S^0\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{44}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0120<sup>+0.0070</sup><sub>-0.0035</sub></b> OUR AVERAGE			

0.022 ±0.016	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts
0.011 ±0.002 <sup>+0.007</sup> <sub>-0.003</sub>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.011 ±0.002 <sup>+0.005</sup> <sub>-0.003</sub>	ASNER	04A CLEO	See MURAMATSU 02
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<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K^*(1680)^-\pi^+, K^*(1680)^-\rightarrow K_S^0\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{45}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.016±0.013</b> OUR AVERAGE			

0.007±0.019	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts
0.022±0.004 <sup>+0.018</sup> <sub>-0.015</sub>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.023±0.005 <sup>+0.007</sup> <sub>-0.014</sub>	ASNER	04A CLEO	See MURAMATSU 02
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<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K^*(892)^+\pi^-, K^*(892)^+\rightarrow K_S^0\pi^+)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{46}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis. This is a doubly Cabibbo-suppressed mode.

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.0<sup>+2.0</sup><sub>-1.2</sub></b> OUR AVERAGE			

4.6±2.3	<sup>1</sup> AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts
3.4±1.3 <sup>+4.1</sup> <sub>-0.4</sub>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.4±1.3 <sup>+3.6</sup> <sub>-0.5</sub>	ASNER	04A CLEO	See MURAMATSU 02
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<sup>1</sup> The error on this AUBERT 08AL value includes both statistical and systematic uncertainties; the latter dominates.

 $\Gamma(K_0^*(1430)^+\pi^-, K_0^*(1430)^+\rightarrow K_S^0\pi^+)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{47}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis. This is a doubly Cabibbo-suppressed mode.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5 × 10<sup>-4</sup></b>	95	AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts

 $\Gamma(K_2^*(1430)^+\pi^-, K_2^*(1430)^+\rightarrow K_S^0\pi^+)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{48}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis. This is a doubly Cabibbo-suppressed mode.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.2 × 10<sup>-3</sup></b>	95	AUBERT	08AL BABR	Dalitz fit, $\approx 487$ k evts

 $\Gamma(K_S^0\pi^+\pi^- \text{ nonresonant})/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{49}/\Gamma_{35}$ 

This is the “fit fraction” from the Dalitz-plot analysis. Neither FRABETTI 94G nor ALBRECHT 93D (quoted in many of the earlier submodes of  $K_S^0\pi^+\pi^-$ ) sees evidence for a nonresonant component.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.009±0.004<sup>+0.020</sup><sub>-0.004</sub></b>	MURAMATSU 02	CLE2	Dalitz fit, 5299 evts

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.007±0.007 <sup>+0.021</sup> <sub>-0.006</sub>	ASNER	04A CLEO	See MURAMATSU 02
0.263±0.024±0.041	ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.26 ±0.08 ±0.05	FRABETTI	92B E687	$\gamma$ Be, $E_\gamma=$ 221 GeV
0.33 ±0.05 ±0.10	ADLER	87 MRK3	$e^+e^-$ 3.77 GeV

 $\Gamma(K^-\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{50}/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.3 ±0.8</b> OUR FIT	Error includes scale factor of 3.1.			

<b>14.956±0.074±0.335</b>	BONVICINI 14	CLEO	All CLEO-c runs
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• • • We do not use the following data for averages, fits, limits, etc. • • •

14.57 ±0.12 ±0.38	<sup>1</sup> DOBBS	07 CLEO	See BONVICINI 14
14.9 ±0.3 ±0.5	<sup>1</sup> HE	05 CLEO	See DOBBS 07
13.3 ±1.2 ±1.3	ADLER	88c MRK3	$e^+e^-$ 3.77 GeV
11.7 ±4.3	<sup>2</sup> SCHINDLER	81 MRK2	$e^+e^-$ 3.771 GeV

<sup>1</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.

<sup>2</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.68 \pm 0.23$  nb. We use the MARK-3 (ADLER 88c) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

 $\Gamma(K^-\pi^+\pi^0)/\Gamma(K^-\pi^+)$   $\Gamma_{50}/\Gamma_{31}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.63±0.22</b> OUR FIT	Error includes scale factor of 3.0.			

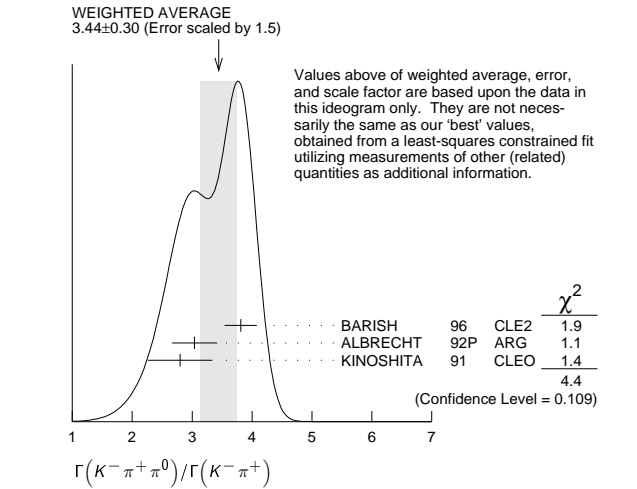
**3.44±0.30** OUR AVERAGE Error includes scale factor of 1.5. See the ideogram below.

3.81±0.07±0.26	10k	BARISH	96 CLE2	$e^+e^- \approx 7(4S)$
3.04±0.16±0.34	931	<sup>1</sup> ALBRECHT	92P ARG	$e^+e^- \approx 10$ GeV
2.8 ±0.14±0.52	1050	KINOSHITA	91 CLEO	$e^+e^- \sim 10.7$ GeV

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

# Meson Particle Listings

$D^0$



**$\Gamma(K^- \rho^+) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{51} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.78 ± 0.04 OUR AVERAGE</b>			
0.788 ± 0.019 ± 0.048	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts
0.765 ± 0.041 ± 0.054	FRABETTI 94G	E687	Dalitz fit, 530 evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.647 ± 0.039 ± 0.150	ANJOS	93	E691 $\gamma$ Be 90–260 GeV
0.81 ± 0.03 ± 0.06	ADLER	87	MRK3 $e^+ e^-$ 3.77 GeV

**$\Gamma(K^- \rho(1700)^+ , \rho(1700)^+ \rightarrow \pi^+ \pi^0) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{52} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.057 ± 0.008 ± 0.009</b>	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts

**$\Gamma(K^*(892)^- \pi^+ , K^*(892)^- \rightarrow K^- \pi^0) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{53} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.160 ± 0.025 OUR AVERAGE</b> <b>−0.013</b>			
0.161 ± 0.007 ± 0.027 −0.011	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts
0.148 ± 0.028 ± 0.049	FRABETTI 94G	E687	Dalitz fit, 530 evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.084 ± 0.011 ± 0.012	ANJOS	93	E691 $\gamma$ Be 90–260 GeV
0.12 ± 0.02 ± 0.03	ADLER	87	MRK3 $e^+ e^-$ 3.77 GeV

**$\Gamma(K^*(892)^0 \pi^0 , K^*(892)^0 \rightarrow K^- \pi^+) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{54} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.135 ± 0.016 OUR AVERAGE</b>			
0.127 ± 0.009 ± 0.016	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts
0.165 ± 0.031 ± 0.015	FRABETTI 94G	E687	Dalitz fit, 530 evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.142 ± 0.018 ± 0.024	ANJOS	93	E691 $\gamma$ Be 90–260 GeV
0.13 ± 0.02 ± 0.03	ADLER	87	MRK3 $e^+ e^-$ 3.77 GeV

**$\Gamma(K_0^*(1430)^- \pi^+ , K_0^*(1430)^- \rightarrow K^- \pi^0) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{55} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.033 ± 0.006 ± 0.014</b>	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts

**$\Gamma(K_0^*(1430)^0 \pi^0 , K_0^*(1430)^0 \rightarrow K^- \pi^+) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{56} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.041 ± 0.006 ± 0.032</b> <b>−0.009</b>	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts

**$\Gamma(K^*(1680)^- \pi^+ , K^*(1680)^- \rightarrow K^- \pi^0) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{57} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.013 ± 0.003 ± 0.004</b>	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts

**$\Gamma(K^- \pi^+ \pi^0 \text{ nonresonant}) / \Gamma(K^- \pi^+ \pi^0)$**   **$\Gamma_{58} / \Gamma_{50}$**

This is the "fit fraction" from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.080 ± 0.040</b> <b>−0.014</b>			
0.075 ± 0.009 ± 0.056 −0.011	KOPP	01	CLE2 Dalitz fit, ≈ 7,000 evts
0.101 ± 0.033 ± 0.040	FRABETTI 94G	E687	Dalitz fit, 530 evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.036 ± 0.004 ± 0.018	ANJOS	93	E691 $\gamma$ Be 90–260 GeV
0.09 ± 0.02 ± 0.04	ADLER	87	MRK3 $e^+ e^-$ 3.77 GeV
0.51 ± 0.22	SUMMERS 21	E691	Photoproduction

**$\Gamma(K_S^0 2\pi^0) / \Gamma_{\text{total}}$**   **$\Gamma_{59} / \Gamma$**

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.1 ± 1.1 OUR AVERAGE</b>				Error includes scale factor of 2.2.
10.58 ± 0.38 ± 0.73	1259	LOWREY	11	CLEO $e^+ e^- \approx 3.77$ GeV
8.34 ± 0.45 ± 0.42		ASNER	08	CLEO $e^+ e^- \rightarrow D^0 \bar{D}^0$ , 3.77 GeV

**$\Gamma(K_S^0 (2\pi^0)\text{-S-wave}) / \Gamma(K_S^0 2\pi^0)$**   **$\Gamma_{60} / \Gamma_{59}$**

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>28.9 ± 6.3 ± 3.1</b>	LOWREY	11	CLEO Dalitz analysis, 1259 evts

**$\Gamma(K^*(892)^0 \pi^0 , K^*(892)^0 \rightarrow K_S^0 \pi^0) / \Gamma(K_S^0 \pi^0)$**   **$\Gamma_{61} / \Gamma_{33}$**

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>65.6 ± 5.3 ± 2.5</b>	LOWREY	11	CLEO Dalitz analysis, 1259 evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
55 $^{+13}_{-10}$ ± 7	PROCARIO	93b	CLE2 Dalitz plot fit, 122 evts

**$\Gamma(K^*(1430)^0 \pi^0 , K^*(1430)^0 \rightarrow K_S^0 \pi^0) / \Gamma(K_S^0 \pi^0)$**   **$\Gamma_{62} / \Gamma_{59}$**

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.49 ± 0.45 ± 2.51</b>	LOWREY	11	CLEO Dalitz analysis, 1259 evts

**$\Gamma(K^*(1680)^0 \pi^0 , K^*(1680)^0 \rightarrow K_S^0 \pi^0) / \Gamma(K_S^0 2\pi^0)$**   **$\Gamma_{63} / \Gamma_{59}$**

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>11.2 ± 2.7 ± 2.5</b>	LOWREY	11	CLEO Dalitz analysis, 1259 evts

**$\Gamma(K_S^0 f_2(1270) , f_2 \rightarrow 2\pi^0) / \Gamma(K_S^0 2\pi^0)$**   **$\Gamma_{64} / \Gamma_{59}$**

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>2.48 ± 0.91 ± 0.78</b>	LOWREY	11	CLEO Dalitz analysis, 1259 evts

**$\Gamma(2K_S^0 , \text{one } K_S^0 \rightarrow 2\pi^0) / \Gamma(K_S^0 2\pi^0)$**   **$\Gamma_{65} / \Gamma_{59}$**

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>3.46 ± 0.92 ± 0.66</b>	LOWREY	11	CLEO Dalitz analysis, 1259 evts

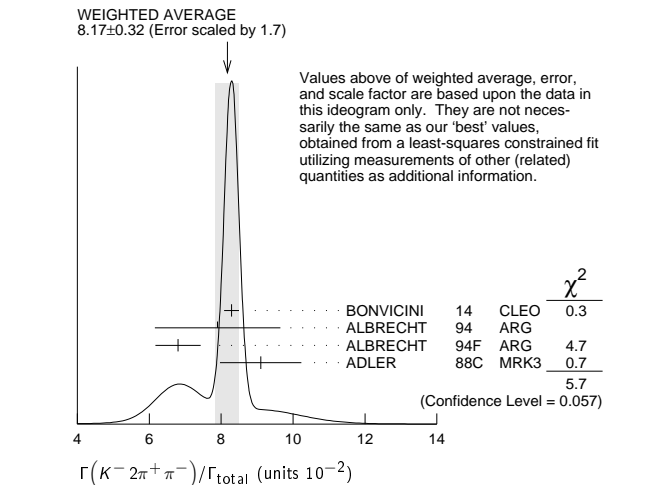
**$\Gamma(K_S^0 2\pi^0 \text{ nonresonant}) / \Gamma(K_S^0 \pi^0)$**   **$\Gamma_{66} / \Gamma_{33}$**

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.37 ± 0.08 ± 0.04	PROCARIO	93b	CLE2 Dalitz plot fit, 122 evts

**$\Gamma(K^- 2\pi^+ \pi^-) / \Gamma_{\text{total}}$**   **$\Gamma_{67} / \Gamma$**

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.06 ± 0.23 OUR FIT</b>				Error includes scale factor of 1.5.
<b>8.17 ± 0.32 OUR AVERAGE</b>				Error includes scale factor of 1.7. See the ideogram below.
8.287 ± 0.043 ± 0.200		BONVICINI	14	CLEO All CLEO-c runs
7.9 ± 1.5 ± 0.9		<sup>1</sup> ALBRECHT	94	ARG $e^+ e^- \approx \Upsilon(4S)$
6.80 ± 0.27 ± 0.57	1.4k	<sup>2</sup> ALBRECHT	94F	ARG $e^+ e^- \approx \Upsilon(4S)$
9.1 ± 0.8 ± 0.8	992	ADLER	88c	MRK3 $e^+ e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.30 ± 0.07 ± 0.20		<sup>3</sup> DOBBS	07	CLEO See BONVICINI 14
8.3 ± 0.2 ± 0.3	15k	<sup>3</sup> HE	05	CLEO See DOBBS 07
11.7 ± 2.5	185	<sup>4</sup> SCHINDLER	81	MRK2 $e^+ e^-$ 3.77 GeV
6.2 ± 1.9	44	<sup>5</sup> PERUZZI	77	LGW $e^+ e^-$ 3.77 GeV

- <sup>1</sup> ALBRECHT 94 uses  $D^0$  mesons from  $\bar{B}^0 \rightarrow D^{*+} \ell^- \bar{\nu}_\ell$  decays. This is a different set of events than used by ALBRECHT 94F.
- <sup>2</sup> See the footnote on the ALBRECHT 94F measurement of  $\Gamma(K^- \pi^+) / \Gamma_{\text{total}}$  for the method used.
- <sup>3</sup> DOBBS 07 and HE 05 use single- and double-tagged events in an overall fit. DOBBS 07 supersedes HE 05.
- <sup>4</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be 0.68 ± 0.11 nb. We use the MARK-3 (ADLER 88c) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.
- <sup>5</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be 0.36 ± 0.10 nb. We use the MARK-3 (ADLER 88c) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.



$\Gamma(K^- 2\pi^+ \pi^-)/\Gamma(K^- \pi^+)$   $\Gamma_{67}/\Gamma_{31}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.05±0.06 OUR FIT</b> Error includes scale factor of 1.4.				
<b>1.97±0.09 OUR AVERAGE</b>				
1.94±0.07+0.09 -0.11		JUN	00	SELX $\Sigma^-$ nucleus, 600 GeV
1.7 ± 0.2 ± 0.2	1745	ANJOS	92c	E691 $\gamma$ Be 90–260 GeV
1.90±0.25±0.20	337	ALVAREZ	91B	NA14 Photoproduction
2.12±0.16±0.09		BORTOLETTO	088	CLEO $e^+e^-$ 10.55 GeV
2.17±0.28±0.23		ALBRECHT	85F	ARG $e^+e^-$ 10 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.0 ± 0.9	48	BAILEY	86	ACCM $\pi^-$ Be fixed target
2.0 ± 1.0	10	BAILEY	83B	SPEC $\pi^-$ Be $\rightarrow D^0$
2.2 ± 0.8	214	PICCOLO	77	MRK1 $e^+e^-$ 4.03, 4.41 GeV

 $\Gamma(K^- \pi^+ \rho^0 \text{total})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{68}/\Gamma_{67}$ 

This includes  $K^- a_1(1260)^+$ ,  $\bar{K}^*(892)^0 \rho^0$ , etc. The next entry gives the specifically 3-body fraction. We rely on the MARKIII and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.835±0.035 OUR AVERAGE</b>			
0.80 ± 0.03 ± 0.05	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts
0.855±0.032±0.030	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.98 ± 0.12 ± 0.10	ALVAREZ	91B	NA14 Photoproduction

 $\Gamma(K^- \pi^+ \rho^0 \text{3-body})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{69}/\Gamma_{67}$ 

We rely on the MARKIII and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.063±0.028 OUR AVERAGE</b>				
0.05 ± 0.03 ± 0.02	ANJOS	92c	E691	1745 $K^- 2\pi^+ \pi^-$ evts
0.084±0.022±0.04	COFFMAN	92B	MRK3	1281 ± 45 $K^- 2\pi^+ \pi^-$ evts
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.77 ± 0.06 ± 0.06	1	ALVAREZ	91B	NA14 Photoproduction
0.85 +0.11 -0.22	180	PICCOLO	77	MRK1 $e^+e^-$ 4.03, 4.41 GeV

<sup>1</sup> This value is for  $\rho^0(K^- \pi^+)$ -nonresonant. ALVAREZ 91B cannot determine what fraction of this is  $K^- a_1(1260)^+$ .

 $\Gamma(\bar{K}^*(892)^0 \rho^0)/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{101}/\Gamma_{67}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included. We rely on the MARKIII and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.195±0.03±0.03</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.34 ± 0.09 ± 0.09		ALVAREZ	91B	NA14 Photoproduction
0.75 ± 0.3	5	BAILEY	83B	SPEC $\pi$ Be $\rightarrow D^0$
0.15 +0.16 -0.15	20	PICCOLO	77	MRK1 $e^+e^-$ 4.03, 4.41 GeV

 $\Gamma(\bar{K}^*(892)^0 \rho^0 \text{transverse})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{102}/\Gamma_{67}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.213±0.024±0.075</b>			
	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 S\text{-wave})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{103}/\Gamma_{67}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.375±0.045±0.06</b>			
	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 \rho^0 S\text{-wave long.})/\Gamma_{\text{total}}$   $\Gamma_{104}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.003</b>				
	90	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 \rho^0 P\text{-wave})/\Gamma_{\text{total}}$   $\Gamma_{105}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.003</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.009	90	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 \rho^0 D\text{-wave})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{106}/\Gamma_{67}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.255±0.045±0.06</b>			
	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K^- \pi^+ f_0(980))/\Gamma_{\text{total}}$   $\Gamma_{107}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.011	90	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 f_0(980))/\Gamma_{\text{total}}$   $\Gamma_{108}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.007	90	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K^- a_1(1260)^+)/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{97}/\Gamma_{67}$ 

Unseen decay modes of the  $a_1(1260)^+$  are included, assuming that the  $a_1(1260)^+$  decays entirely to  $\rho\pi$  [or at least to  $(\pi\pi)_{I=1}\pi$ ].

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.97 ± 0.14 OUR AVERAGE</b>			
0.94 ± 0.13 ± 0.20	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts
0.984±0.048±0.16	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K^- a_2(1320)^+)/\Gamma_{\text{total}}$   $\Gamma_{98}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.002</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.006	90	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K_1(1270)^- \pi^+)/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{109}/\Gamma_{67}$ 

Unseen decay modes of the  $K_1(1270)^-$  are included. The MARK3 and E691 experiments disagree considerably here.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.194±0.056±0.088</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.013	90	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K_1(1400)^- \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{110}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.012</b>				
	90	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(1410)^- \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{111}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.012	90	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \text{total})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{99}/\Gamma_{67}$ 

This includes  $\bar{K}^*(892)^0 \rho^0$ , etc. The next entry gives the specifically 3-body fraction. Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.30±0.06±0.03</b>			
	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \text{3-body})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{100}/\Gamma_{67}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.18 ± 0.04 OUR AVERAGE</b>			
0.165±0.03 ± 0.045	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts
0.210±0.027±0.06	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K^- 2\pi^+ \pi^- \text{nonresonant})/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{75}/\Gamma_{67}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.233±0.032 OUR AVERAGE</b>			
0.23 ± 0.02 ± 0.03	ANJOS	92c	E691 1745 $K^- 2\pi^+ \pi^-$ evts
0.242±0.025±0.06	COFFMAN	92B	MRK3 1281 ± 45 $K^- 2\pi^+ \pi^-$ evts

 $\Gamma(K_S^0 \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{76}/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.2±0.6 OUR FIT</b>				
<b>5.2±1.1±1.2</b>	140	COFFMAN	92B	MRK3 $e^+e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				

6.7+1.6  
-1.7 <sup>1</sup> BARLAG 92c computes the branching fraction using topological normalization.

 $\Gamma(K_S^0 \pi^+ \pi^- \pi^0)/\Gamma(K_S^0 \pi^+ \pi^-)$   $\Gamma_{76}/\Gamma_{35}$ 

Branching fractions for submodes of this mode with narrow resonances (the  $\eta$ ,  $\omega$ ,  $\eta'$ ) are fairly well determined (see below). COFFMAN 92B gives fractions of  $K^*$  and  $\rho$  submodes, but with only 140±28 events above background could not determine them with much accuracy. We omit those measurements here; they are in our 2008 Review (Physics Letters **B667** 1 (2008)).

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.84±0.20 OUR FIT</b>				
<b>1.86±0.23 OUR AVERAGE</b>				
1.80±0.20±0.21	190	<sup>1</sup> ALBRECHT	92P	ARG $e^+e^- \approx$ 10 GeV
2.8 ± 0.8 ± 0.8	46	ANJOS	92c	E691 $\gamma$ Be 90–260 GeV
1.85±0.26±0.30	158	KINOSHITA	91	CLEO $e^+e^- \sim$ 10.7 GeV

<sup>1</sup> This value is calculated from numbers in Table I of ALBRECHT 92P.

 $\Gamma(K_S^0 \eta)/\Gamma_{\text{total}}$   $\Gamma_{94}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.42±0.15±0.28	ASNER	08	CLEO See MENDEZ 10

# Meson Particle Listings

$D^0$

$\Gamma(K_S^0 \eta)/[\Gamma(K^- \pi^+) + \Gamma(K^+ \pi^-)]$   $\Gamma_{94}/(\Gamma_{31} + \Gamma_{222})$   
Unseen decay modes of the  $\eta$  are included.

VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>12.3 \pm 0.8</math> OUR FIT</b>				
<b><math>12.3 \pm 0.3 \pm 0.7</math></b>	2864 $\pm$ 65	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV

$\Gamma(K_S^0 \eta)/\Gamma(K_S^0 \pi^+ \pi^-)$   $\Gamma_{94}/\Gamma_{33}$   
Unseen decay modes of the  $\eta$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.32 \pm 0.04 \pm 0.03$	225 $\pm$ 30	PROCARIO	93B	CLE2 $\eta \rightarrow \gamma \gamma$

$\Gamma(K_S^0 \eta)/\Gamma(K_S^0 \pi^+ \pi^-)$   $\Gamma_{94}/\Gamma_{35}$   
Unseen decay modes of the  $\eta$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.14 \pm 0.02 \pm 0.02$	80 $\pm$ 12	PROCARIO	93B	CLE2 $\eta \rightarrow \pi^+ \pi^- \pi^0$

$\Gamma(K_S^0 \omega)/\Gamma_{\text{total}}$   $\Gamma_{95}/\Gamma$   
Unseen decay modes of the  $\omega$  are included.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>1.11 \pm 0.06</math> OUR FIT</b>			
<b><math>1.12 \pm 0.04 \pm 0.05</math></b>	ASNER 08	CLEO	$e^+ e^- \rightarrow D^0 \bar{D}^0$ , 3.77 GeV

$\Gamma(K_S^0 \omega)/\Gamma(K^- \pi^+)$   $\Gamma_{95}/\Gamma_{31}$   
Unseen decay modes of the  $\omega$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
$0.50 \pm 0.18 \pm 0.10$	ALBRECHT 89D	ARG	$e^+ e^-$ 10 GeV

$\Gamma(K_S^0 \omega)/\Gamma(K_S^0 \pi^+ \pi^-)$   $\Gamma_{95}/\Gamma_{35}$   
Unseen decay modes of the  $\omega$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.389 \pm 0.033</math> OUR FIT</b>	Error includes scale factor of 1.1.			
<b><math>0.33 \pm 0.09</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.			
$0.29 \pm 0.08 \pm 0.05$	16	<sup>1</sup> ALBRECHT	92P	ARG $e^+ e^- \approx 10$ GeV
$0.54 \pm 0.14 \pm 0.16$	40	KINOSHITA	91	CLEO $e^+ e^- \approx 10.7$ GeV

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K_S^0 \omega)/\Gamma(K_S^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{95}/\Gamma_{76}$   
Unseen decay modes of the  $\omega$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.212 \pm 0.026</math> OUR FIT</b>			
<b><math>0.220 \pm 0.048 \pm 0.0116</math></b>	COFFMAN 92B	MRK3	1281 $\pm$ 45 $K^- 2\pi^+ \pi^-$ evts

$\Gamma(K_S^0 \eta'(958))/[\Gamma(K^- \pi^+) + \Gamma(K^+ \pi^-)]$   $\Gamma_{96}/(\Gamma_{31} + \Gamma_{222})$   
Unseen decay modes of the  $\eta'(958)$  are included.

VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>24.1 \pm 1.3</math> OUR FIT</b>				
<b><math>24.3 \pm 0.8 \pm 1.1</math></b>	1321 $\pm$ 42	MENDEZ	10	CLEO $e^+ e^-$ at 3774 MeV

$\Gamma(K_S^0 \eta'(958))/\Gamma(K_S^0 \pi^+ \pi^-)$   $\Gamma_{96}/\Gamma_{35}$   
Unseen decay modes of the  $\eta'(958)$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.332 \pm 0.025</math> OUR FIT</b>				
<b><math>0.32 \pm 0.04</math> OUR AVERAGE</b>				
$0.31 \pm 0.02 \pm 0.04$	594	PROCARIO	93B	CLE2 $\eta' \rightarrow \eta \pi^+ \pi^-, \rho^0 \gamma$
$0.37 \pm 0.13 \pm 0.06$	18	<sup>1</sup> ALBRECHT	92P	ARG $e^+ e^- \approx 10$ GeV

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K^- \pi^+ 2\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{79}/\Gamma$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.177 \pm 0.029$		<sup>1</sup> BARLAG	92c	ACCM $\pi^-$ Cu 230 GeV
$0.149 \pm 0.037 \pm 0.030$	24	<sup>2</sup> ADLER	88c	MRK3 $e^+ e^-$ 3.77 GeV
$0.209 \pm 0.074 \pm 0.043$	9	<sup>1</sup> AGUILAR-...	87F	HYBR $\pi p, pp$ 360, 400 GeV

<sup>1</sup> AGUILAR-BENITEZ 87F and BARLAG 92c compute the branching fraction using topological normalization. They do not distinguish the presence of a third  $\pi^0$ , and thus are not included in the average.

<sup>2</sup> ADLER 88c uses an absolute normalization method finding this decay channel opposite a detected  $\bar{D}^0 \rightarrow K^+ \pi^-$  in pure  $D\bar{D}$  events.

$\Gamma(K^- 2\pi^+ \pi^- \pi^0)/\Gamma(K^- \pi^+)$   $\Gamma_{80}/\Gamma_{31}$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>1.08 \pm 0.10</math> OUR FIT</b>				
<b><math>0.98 \pm 0.11 \pm 0.11</math></b>	225	<sup>1</sup> ALBRECHT	92P	ARG $e^+ e^- \approx 10$ GeV

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K^- 2\pi^+ \pi^- \pi^0)/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{80}/\Gamma_{67}$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.53 \pm 0.05</math> OUR FIT</b>				
<b><math>0.56 \pm 0.07</math> OUR AVERAGE</b>				
$0.55 \pm 0.07 \pm 0.12 \pm 0.09$	167	KINOSHITA	91	CLEO $e^+ e^- \sim 10.7$ GeV
$0.57 \pm 0.06 \pm 0.05$	180	ANJOS	90D	E691 Photoproduction

$\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0)/\Gamma(K^- 2\pi^+ \pi^- \pi^0)$   $\Gamma_{112}/\Gamma_{80}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.45 \pm 0.15 \pm 0.15</math></b>	ANJOS 90D	E691	Photoproduction

$\Gamma(\bar{K}^*(892)^0 \eta)/\Gamma(K^- \pi^+)$   $\Gamma_{113}/\Gamma_{31}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\eta$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.58 \pm 0.19 \pm 0.24 \pm 0.28$	46	KINOSHITA	91	CLEO $e^+ e^- \sim 10.7$ GeV

$\Gamma(\bar{K}^*(892)^0 \eta)/\Gamma(K^- \pi^+ \pi^0)$   $\Gamma_{113}/\Gamma_{50}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\eta$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.13 \pm 0.02 \pm 0.03$	214	PROCARIO	93B	CLE2 $\bar{K}^* 0 \eta \rightarrow K^- \pi^+ / \gamma \gamma$

$\Gamma(K_S^0 \eta \pi^0)/\Gamma(K_S^0 \pi^0)$   $\Gamma_{84}/\Gamma_{33}$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.46 \pm 0.07 \pm 0.06</math></b>	155 $\pm$ 22	<sup>1</sup> RUBIN	04	CLEO $e^+ e^- \approx 10$ GeV

<sup>1</sup> The  $\eta$  here is detected in its  $\gamma \gamma$  mode, but other  $\eta$  modes are included in the value given.

$\Gamma(K_S^0 a_0(980), a_0(980) \rightarrow \eta \pi^0)/\Gamma(K_S^0 \eta \pi^0)$   $\Gamma_{85}/\Gamma_{84}$   
This is the “fit fraction” from the Dalitz-plot analysis, with interference.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.19 \pm 0.09 \pm 0.26</math></b>	<sup>1</sup> RUBIN 04	CLEO	Dalitz fit, 155 evts

<sup>1</sup> In addition to  $K_S^0 a_0(980)$  and  $\bar{K}^*(892)^0 \eta$  modes, RUBIN 04 finds a fit fraction of  $0.246 \pm 0.092 \pm 0.091$  for other, undetermined modes.

$\Gamma(\bar{K}^*(892)^0 \eta, \bar{K}^*(892)^0 \rightarrow K_S^0 \pi^0)/\Gamma(K_S^0 \eta \pi^0)$   $\Gamma_{86}/\Gamma_{84}$   
This is the “fit fraction” from the Dalitz-plot analysis, with interference.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.293 \pm 0.062 \pm 0.035</math></b>	<sup>1</sup> RUBIN 04	CLEO	Dalitz fit, 155 evts

<sup>1</sup> See the note on RUBIN 04 in the preceding data block.

$\Gamma(K^- \pi^+ \omega)/\Gamma(K^- \pi^+)$   $\Gamma_{114}/\Gamma_{31}$   
Unseen decay modes of the  $\omega$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.78 \pm 0.12 \pm 0.10</math></b>	99	<sup>1</sup> ALBRECHT	92P	ARG $e^+ e^- \approx 10$ GeV

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^*(892)^0 \omega)/\Gamma(K^- \pi^+)$   $\Gamma_{115}/\Gamma_{31}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\omega$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.28 \pm 0.11 \pm 0.04</math></b>	17	<sup>1</sup> ALBRECHT	92P	ARG $e^+ e^- \approx 10$ GeV

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K^- \pi^+ \eta'(958))/\Gamma(K^- 2\pi^+ \pi^-)$   $\Gamma_{116}/\Gamma_{67}$   
Unseen decay modes of the  $\eta'(958)$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.093 \pm 0.014 \pm 0.019</math></b>	286	PROCARIO	93B	CLE2 $\eta' \rightarrow \eta \pi^+ \pi^-, \rho^0 \gamma$

$\Gamma(\bar{K}^*(892)^0 \eta'(958))/\Gamma(K^- \pi^+ \eta'(958))$   $\Gamma_{117}/\Gamma_{116}$   
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.15</math></b>	90	PROCARIO	93B	CLE2

$\Gamma(K_S^0 2\pi^+ 2\pi^-)/\Gamma(K_S^0 \pi^+ \pi^-)$   $\Gamma_{87}/\Gamma_{35}$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.095 \pm 0.005 \pm 0.007</math></b>	1283 $\pm$ 57	LINK	04D	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.07 \pm 0.02 \pm 0.01$	11	<sup>1</sup> ALBRECHT	92P	ARG $e^+ e^- \approx 10$ GeV
$0.149 \pm 0.026$	56	ANMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
$0.18 \pm 0.07 \pm 0.04$	6	ANJOS	90D	E691 Photoproduction

<sup>1</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K_S^0 \rho^0 \pi^+ \pi^-, \text{no } K^*(892)^-)/\Gamma(K_S^0 2\pi^+ 2\pi^-)$   $\Gamma_{88}/\Gamma_{87}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.40 \pm 0.24 \pm 0.07</math></b>	LINK 04D	FOCS	$\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(K^*(892)^- 2\pi^+ \pi^-, K^*(892)^- \rightarrow K_S^0 \pi^-, \text{no } \rho^0)/\Gamma(K_S^0 2\pi^+ 2\pi^-)$   $\Gamma_{89}/\Gamma_{87}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.17 \pm 0.28 \pm 0.02</math></b>	LINK 04D	FOCS	$\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(K^*(892)^- \rho^0 \pi^+, K^*(892)^- \rightarrow K_S^0 \pi^-)/\Gamma(K_S^0 2\pi^+ 2\pi^-)$   $\Gamma_{90}/\Gamma_{87}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.60 \pm 0.21 \pm 0.09</math></b>	LINK 04D	FOCS	$\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(K_S^0 2\pi^+ 2\pi^- \text{nonresonant})/\Gamma(K_S^0 2\pi^+ 2\pi^-)$   $\Gamma_{91}/\Gamma_{87}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.46</math></b>	90	LINK	04D	FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV



See key on page 601

## Meson Particle Listings

 $D^0$ 

$\Gamma(K^- 3\pi^+ 2\pi^-)/\Gamma(K^- 2\pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{93}/\Gamma_{67}</math></span>				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.70 \pm 0.58 \pm 0.38</math></b>	$48 \pm 10$	LINK	04B FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV

Hadronic modes with three  $K$ 's

$\Gamma(K_S^0 K^+ K^-)/\Gamma(K_S^0 \pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{118}/\Gamma_{35}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.158 \pm 0.001 \pm 0.005</math></b>	$14k \pm 116$	AUBERT,B	05J BABR	$e^+ e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.20 \pm 0.05 \pm 0.04$	47	FRABETTI	92b E687	$\gamma$ Be, $\overline{E}_\gamma \approx 221$ GeV
$0.170 \pm 0.022$	136	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
$0.24 \pm 0.08$		BEBEK	86 CLEO	$e^+ e^-$ near $\gamma(4S)$
$0.185 \pm 0.055$	52	ALBRECHT	85b ARG	$e^+ e^- 10$ GeV

$\Gamma(K_S^0 a_0(980)^0, a_0^0 \rightarrow K^+ K^-)/\Gamma(K_S^0 K^+ K^-)$ <span style="float:right"><math>\Gamma_{119}/\Gamma_{118}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.664 \pm 0.016 \pm 0.070</math></b>	AUBERT,B	05J BABR	Dalitz fit, 12540 $\pm$ 112 evts	

$\Gamma(K^- a_0(980)^+, a_0^+ \rightarrow K^+ K_S^0)/\Gamma(K_S^0 K^+ K^-)$ <span style="float:right"><math>\Gamma_{120}/\Gamma_{118}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.134 \pm 0.011 \pm 0.037</math></b>	AUBERT,B	05J BABR	Dalitz fit, 12540 $\pm$ 112 evts	

$\Gamma(K^+ a_0(980)^-, a_0^- \rightarrow K^- K_S^0)/\Gamma(K_S^0 K^+ K^-)$ <span style="float:right"><math>\Gamma_{121}/\Gamma_{118}</math></span>				
This is a doubly Cabibbo-suppressed mode.				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.025</math></b>	95	AUBERT,B	05J BABR	Dalitz fit, 12540 $\pm$ 112 evts

$\Gamma(K_S^0 f_0(980), f_0 \rightarrow K^+ K^-)/\Gamma(K_S^0 K^+ K^-)$ <span style="float:right"><math>\Gamma_{122}/\Gamma_{118}</math></span>				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.021</math></b>	95	AUBERT,B	05J BABR	Dalitz fit, 12540 $\pm$ 112 evts

$\Gamma(K_S^0 \phi, \phi \rightarrow K^+ K^-)/\Gamma(K_S^0 K^+ K^-)$ <span style="float:right"><math>\Gamma_{123}/\Gamma_{118}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.459 \pm 0.007 \pm 0.007</math></b>	AUBERT,B	05J BABR	Dalitz fit, 12540 $\pm$ 112 evts	

$\Gamma(K_S^0 f_0(1370), f_0 \rightarrow K^+ K^-)/\Gamma(K_S^0 K^+ K^-)$ <span style="float:right"><math>\Gamma_{124}/\Gamma_{118}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.038 \pm 0.007 \pm 0.023</math></b>	<sup>1</sup> AUBERT,B	05J BABR	Dalitz fit, 12540 $\pm$ 112 evts	
<sup>1</sup> AUBERT,B 05J calls the mode $K_S^0 f_0(1400)$ , but insofar as it is seen here at all, it is certainly the same as $f_0(1370)$ .				

$\Gamma(3K_S^0)/\Gamma(K_S^0 \pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{125}/\Gamma_{35}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.2 \pm 0.4</math> OUR AVERAGE</b>				
$3.58 \pm 0.54 \pm 0.52$	$170 \pm 26$	LINK	05A FOCUS	$\gamma$ Be, $\overline{E}_\gamma \approx 180$ GeV
$2.78 \pm 0.38 \pm 0.48$	61	ASNER	96b CLE2	$e^+ e^- \approx \gamma(4S)$
$7.0 \pm 2.4 \pm 1.2$	$10 \pm 3$	FRABETTI	94J E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
$3.2 \pm 1.0$	22	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
$3.4 \pm 1.4 \pm 1.0$	5	ALBRECHT	90c ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(K^+ 2K^- \pi^+)/\Gamma(K^- 2\pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{126}/\Gamma_{67}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.0027 \pm 0.0004</math> OUR AVERAGE</b>				Error includes scale factor of 1.1.
$0.00257 \pm 0.00034 \pm 0.00024$	143	LINK	03G FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV
$0.0054 \pm 0.0016 \pm 0.0008$	18	AITALA	01d E791	$\pi^- A, 500$ GeV
$0.0028 \pm 0.0007 \pm 0.0001$	20	FRABETTI	95c E687	$\gamma$ Be, $\overline{E}_\gamma \approx 200$ GeV

$\Gamma(\phi \overline{K}^*(892)^0, \phi \rightarrow K^+ K^-, \overline{K}^*(892)^0 \rightarrow K^- \pi^+)/\Gamma(K^+ 2K^- \pi^+)$ <span style="float:right"><math>\Gamma_{129}/\Gamma_{126}</math></span>				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.48 \pm 0.06 \pm 0.01</math></b>	LINK	03G FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV	

$\Gamma(K^- \pi^+ \phi, \phi \rightarrow K^+ K^-)/\Gamma(K^+ 2K^- \pi^+)$ <span style="float:right"><math>\Gamma_{128}/\Gamma_{126}</math></span>				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.18 \pm 0.06 \pm 0.04</math></b>	LINK	03G FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV	

$\Gamma(K^+ K^- \overline{K}^*(892)^0, \overline{K}^*(892)^0 \rightarrow K^- \pi^+)/\Gamma(K^+ 2K^- \pi^+)$ <span style="float:right"><math>\Gamma_{127}/\Gamma_{126}</math></span>				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.20 \pm 0.07 \pm 0.02</math></b>	LINK	03G FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV	

$\Gamma(K^+ 2K^- \pi^+ \text{nonresonant})/\Gamma(K^+ 2K^- \pi^+)$ <span style="float:right"><math>\Gamma_{130}/\Gamma_{126}</math></span>				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.15 \pm 0.06 \pm 0.02</math></b>	LINK	03G FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV	

$\Gamma(2K_S^0 K^\pm \pi^\mp)/\Gamma(K_S^0 \pi^+ \pi^-)$ <span style="float:right"><math>\Gamma_{131}/\Gamma_{35}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.12 \pm 0.38 \pm 0.20</math></b>	$57 \pm 10$	LINK	05A FOCUS	$\gamma$ Be, $\overline{E}_\gamma \approx 180$ GeV

## Pionic modes

$\Gamma(\pi^+ \pi^-)/\Gamma(K^- \pi^+)$ <span style="float:right"><math>\Gamma_{132}/\Gamma_{31}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.62 \pm 0.05</math> OUR FIT</b>				
<b><math>3.59 \pm 0.06</math> OUR AVERAGE</b>				
$3.594 \pm 0.054 \pm 0.040$	$7334 \pm 97$	ACOSTA	05c CDF	$p\overline{p}, \sqrt{s} = 1.96$ TeV
$3.53 \pm 0.12 \pm 0.06$	3453	LINK	03 FOCUS	$\gamma A, \overline{E}_\gamma \approx 180$ GeV
$3.51 \pm 0.16 \pm 0.17$	710	CSORNA	02 CLE2	$e^+ e^- \approx \gamma(4S)$
$4.0 \pm 0.2 \pm 0.3$	2043	AITALA	98c E791	$\pi^- A, 500$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$3.62 \pm 0.10 \pm 0.08$	$2085 \pm 54$	RUBIN	06 CLEO	See MENDEZ 10
$3.4 \pm 0.7 \pm 0.1$	$76 \pm 15$	ABLIKIM	05f BES	$e^+ e^- \approx \psi(3770)$
$4.3 \pm 0.7 \pm 0.3$	177	FRABETTI	94c E687	$\gamma$ Be $\overline{E}_\gamma \approx 220$ GeV
$3.48 \pm 0.30 \pm 0.23$	227	SELEN	93 CLE2	$e^+ e^- \approx \gamma(4S)$
$5.5 \pm 0.8 \pm 0.5$	120	ANJOS	91d E691	Photoproduction
$5.0 \pm 0.7 \pm 0.5$	110	ALEXANDER	90 CLEO	$e^+ e^- 10.5\text{--}11$ GeV

$\Gamma(\pi^+ \pi^-)/[\Gamma(K^- \pi^+) + \Gamma(K^+ \pi^-)]$ <span style="float:right"><math>\Gamma_{132}/(\Gamma_{31} + \Gamma_{222})</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.60 \pm 0.05</math> OUR FIT</b>				
<b><math>3.70 \pm 0.06 \pm 0.09</math></b>	$6210 \pm 93$	MENDEZ	10 CLEO	$e^+ e^-$ at 3774 MeV

$\Gamma(2\pi^0)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{133}/\Gamma</math></span>				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>8.25 \pm 0.25</math> OUR FIT</b>				
<b><math>8.29 \pm 0.30</math> OUR AVERAGE</b>				
$8.24 \pm 0.21 \pm 0.30$	6k	ABLIKIM	15f BES3	$e^+ e^-$ at 3.773 GeV
$8.4 \pm 0.1 \pm 0.5$	26k	LEES	12L BABR	$e^+ e^- \approx 10.58$ GeV

$\Gamma(2\pi^0)/\Gamma(K^- \pi^+)$ <span style="float:right"><math>\Gamma_{133}/\Gamma_{31}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.05 \pm 0.13 \pm 0.16$	$499 \pm 32$	RUBIN	06 CLEO	See MENDEZ 10
$2.2 \pm 0.4 \pm 0.4$	40	SELEN	93 CLE2	$e^+ e^- \rightarrow \gamma(4S)$

$\Gamma(2\pi^0)/[\Gamma(K^- \pi^+) + \Gamma(K^+ \pi^-)]$ <span style="float:right"><math>\Gamma_{133}/(\Gamma_{31} + \Gamma_{222})</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.09 \pm 0.07</math> OUR FIT</b>				
<b><math>2.06 \pm 0.07 \pm 0.10</math></b>	$1567 \pm 54$	MENDEZ	10 CLEO	$e^+ e^-$ at 3774 MeV

$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma(K^- \pi^+)$ <span style="float:right"><math>\Gamma_{134}/\Gamma_{31}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>37.5 \pm 2.3</math> OUR FIT</b>				Error includes scale factor of 3.0.
<b><math>34.4 \pm 0.5 \pm 1.2</math></b>	$11k \pm 164$	RUBIN	06 CLEO	$e^+ e^-$ at $\psi(3770)$

$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma(K^- \pi^+ \pi^0)$ <span style="float:right"><math>\Gamma_{134}/\Gamma_{50}</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>10.33 \pm 0.25</math> OUR FIT</b>				Error includes scale factor of 2.3.
<b><math>10.41 \pm 0.23</math> OUR AVERAGE</b>				Error includes scale factor of 2.0.
$10.12 \pm 0.04 \pm 0.18$	$123k \pm 490$	ARINSTEIN	08 BELL	$e^+ e^- \approx \gamma(4S)$
$10.59 \pm 0.06 \pm 0.13$	$60k \pm 343$	AUBERT,B	06x BABR	$e^+ e^- \approx \gamma(4S)$

$\Gamma(\rho^+ \pi^-)/\Gamma(\pi^+ \pi^- \pi^0)$ <span style="float:right"><math>\Gamma_{135}/\Gamma_{134}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference. See GASPERO 08 and BHATTACHARYA 10a for isospin decompositions of the $D^0 \rightarrow \pi^+ \pi^0 \pi^-$ Dalitz plot, both based on the amplitudes of AUBERT 07Bj. They quantify the conclusion that the final state is dominantly isospin 0.				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	

<b><math>68.1 \pm 0.6</math> OUR AVERAGE</b>				
$67.8 \pm 0.0 \pm 0.6$	AUBERT	07Bj BABR	Dalitz fit, 45k events	
$76.3 \pm 1.9 \pm 2.5$	CRONIN-HEN..05	CLEO	$e^+ e^- \approx 10$ GeV	

$\Gamma(\rho^0 \pi^0)/\Gamma(\pi^+ \pi^- \pi^0)$ <span style="float:right"><math>\Gamma_{136}/\Gamma_{134}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference.				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>25.9 \pm 1.1</math> OUR AVERAGE</b>				
$26.2 \pm 0.5 \pm 1.1$	AUBERT	07Bj BABR	Dalitz fit, 45k events	
$24.4 \pm 2.0 \pm 2.1$	CRONIN-HEN..05	CLEO	$e^+ e^- \approx 10$ GeV	

$\Gamma(\rho^- \pi^+)/\Gamma(\pi^+ \pi^- \pi^0)$ <span style="float:right"><math>\Gamma_{137}/\Gamma_{134}</math></span>				
This is the "fit fraction" from the Dalitz-plot analysis, with interference.				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>34.6 \pm 0.8</math> OUR AVERAGE</b>				
$34.6 \pm 0.8 \pm 0.3$	AUBERT	07Bj BABR	Dalitz fit, 45k events	
$34.5 \pm 2.4 \pm 1.3$	CRONIN-HEN..05	CLEO	$e^+ e^- \approx 10$ GeV	

$\Gamma(\rho(1450)^+ \pi^-, \rho(1450)^+ \rightarrow \pi^+ \pi^0)/\Gamma(\pi^+ \pi^- \pi^0)$ <span style="float:right"><math>\Gamma_{138}/\Gamma_{134}</math></span>				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.11 \pm 0.07 \pm 0.12</math></b>	AUBERT	07Bj BABR	Dalitz fit, 45k events	

$\Gamma(\rho(1450)^0 \pi^0, \rho(1450)^0 \rightarrow \pi^+ \pi^-)/\Gamma(\pi^+ \pi^- \pi^0)$ <span style="float:right"><math>\Gamma_{139}/\Gamma_{134}</math></span>				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.30 \pm 0.11 \pm 0.07</math></b>	AUBERT	07Bj BABR	Dalitz fit, 45k events	

# Meson Particle Listings

$D^0$

$\Gamma(\rho(1450)^-\pi^+,\rho(1450)^-\rightarrow\pi^-\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{140}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$1.79\pm 0.22\pm 0.12$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(\rho(1700)^+\pi^-, \rho(1700)^+\rightarrow\pi^+\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{141}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$4.1\pm 0.7\pm 0.7$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(\rho(1700)^0\pi^0,\rho(1700)^0\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{142}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$5.0\pm 0.6\pm 1.0$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(\rho(1700)^-\pi^+,\rho(1700)^-\rightarrow\pi^-\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{143}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$3.2\pm 0.4\pm 0.6$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(f_0(980)\pi^0,f_0(980)\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{144}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u> <u>CL%</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$0.25\pm 0.04\pm 0.04$	AUBERT 07BJ BABR Dalitz fit, 45k events
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<0.026 95 <sup>1</sup> CRONIN-HEN..05 CLEO $e^+e^-\approx 10$ GeV	
<sup>1</sup> The CRONIN-HENNESSY 05 fit here includes, in addition to the three $\rho\pi$ charged states, only the $f_0(980)\pi^0$ mode. See also the next entries for limits obtained in the same way for the $f_0(500)\pi^0$ mode and for an S-wave $\pi^+\pi^-$ parametrized using a $K$ -matrix. Our $\rho\pi$ branching ratios, given above, use the fit with the $K$ -matrix S wave.	
$\Gamma(f_0(500)\pi^0,f_0(500)\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{145}/\Gamma_{134}$
The $f_0(500)$ is the $\sigma$ .	
<u>VALUE (units <math>10^{-2}</math>)</u> <u>CL%</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$0.82\pm 0.10\pm 0.10$	AUBERT 07BJ BABR Dalitz fit, 45k events
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<0.21 95 <sup>1</sup> CRONIN-HEN..05 CLEO $e^+e^-\approx 10$ GeV	
<sup>1</sup> See the note on CRONIN-HENNESSY 05 in the proceeding data block.	
$\Gamma((\pi^+\pi^-)_{S\text{-wave}}\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{146}/\Gamma_{134}$
<u>VALUE</u> <u>CL%</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<0.019 95 <sup>1</sup> CRONIN-HEN..05 CLEO $e^+e^-\approx 10$ GeV	
<sup>1</sup> See the note on CRONIN-HENNESSY 05 two data blocks up.	
$\Gamma(f_0(1370)\pi^0,f_0(1370)\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{147}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$0.37\pm 0.11\pm 0.09$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(f_0(1500)\pi^0,f_0(1500)\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{148}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$0.39\pm 0.08\pm 0.07$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(f_0(1710)\pi^0,f_0(1710)\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{149}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$0.31\pm 0.07\pm 0.08$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(f_2(1270)\pi^0,f_2(1270)\rightarrow\pi^+\pi^-)/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{150}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$1.32\pm 0.08\pm 0.10$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(\pi^+\pi^-\pi^0\text{nonresonant})/\Gamma(\pi^+\pi^-\pi^0)$	$\Gamma_{151}/\Gamma_{134}$
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$0.84\pm 0.21\pm 0.12$	AUBERT 07BJ BABR Dalitz fit, 45k events
$\Gamma(3\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{152}/\Gamma$
<u>VALUE</u> <u>CL%</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>	
$<3.5\times 10^{-4}$ 90	RUBIN 06 CLEO $e^+e^-\text{ at } \psi(3770)$
$\Gamma(2\pi^+2\pi^-)/\Gamma(K^-\pi^+)$	$\Gamma_{153}/\Gamma_{31}$
<u>VALUE (units <math>10^{-2}</math>)</u> <u>EVTS</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$19.0\pm 0.6$ OUR FIT Error includes scale factor of 1.1.	
$19.1\pm 0.4\pm 0.6$ 7331 $\pm$ 130 RUBIN 06 CLEO $e^+e^-\text{ at } \psi(3770)$	
$\Gamma(2\pi^+2\pi^-)/\Gamma(K^-2\pi^+\pi^-)$	$\Gamma_{153}/\Gamma_{67}$
<u>VALUE (units <math>10^{-2}</math>)</u> <u>EVTS</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$9.24\pm 0.23$ OUR FIT Error includes scale factor of 1.1.	
$9.20\pm 0.26$ OUR AVERAGE	
9.14 $\pm$ 0.18 $\pm$ 0.22 6360 $\pm$ 115	LINK 07A FOCS $\gamma\text{Be}, \overline{E}_\gamma\approx 180$ GeV
7.9 $\pm$ 1.8 $\pm$ 0.5 162	ABLIKIM 05F BES $e^+e^-\approx \psi(3770)$
9.5 $\pm$ 0.7 $\pm$ 0.2 814	FRABETTI 95c E687 $\gamma\text{Be}, \overline{E}_\gamma\approx 200$ GeV
10.2 $\pm$ 1.3 345	AMMAR 91 CLEO $e^+e^-\approx 10.5$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
11.5 $\pm$ 2.3 $\pm$ 1.6 64	ADAMOVICH 92 OMEG $\pi^-$ 340 GeV
10.8 $\pm$ 2.4 $\pm$ 0.8 79	FRABETTI 92 E687 $\gamma\text{Be}$
9.6 $\pm$ 1.8 $\pm$ 0.7 66	ANJOS 91 E691 $\gamma\text{Be}$ 80–240 GeV

$\Gamma(a_1(1260)^+\pi^-,a_1^+\rightarrow 2\pi^+\pi^-\text{total})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{154}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$60.0\pm 3.0\pm 2.4$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(a_1(1260)^+\pi^-,a_1^+\rightarrow\rho^0\pi^+ \text{ S-wave})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{155}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$43.3\pm 2.5\pm 1.9$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(a_1(1260)^+\pi^-,a_1^+\rightarrow\rho^0\pi^+ \text{ D-wave})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{156}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$2.5\pm 0.5\pm 0.4$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(a_1(1260)^+\pi^-,a_1^+\rightarrow\sigma\pi^+)/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{157}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$8.3\pm 0.7\pm 0.6$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(2\rho^0\text{total})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{158}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$24.5\pm 1.3\pm 1.0$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(2\rho^0,\text{parallel helicities})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{159}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$1.1\pm 0.3\pm 0.3$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(2\rho^0,\text{perpendicular helicities})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{160}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$6.4\pm 0.6\pm 0.5$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(2\rho^0,\text{longitudinal helicities})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{161}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$16.8\pm 1.0\pm 0.8$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(\text{Resonant }(\pi^+\pi^-)\pi^+\pi^-\text{ 3-body total})/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{162}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$20.0\pm 1.2\pm 1.0$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(\sigma\pi^+\pi^-)/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{163}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$8.2\pm 0.9\pm 0.7$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(f_0(980)\pi^+\pi^-,f_0\rightarrow\pi^+\pi^-)/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{164}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$2.4\pm 0.5\pm 0.4$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(f_2(1270)\pi^+\pi^-,f_2\rightarrow\pi^+\pi^-)/\Gamma(2\pi^+2\pi^-)$	$\Gamma_{165}/\Gamma_{153}$
This is the fit fraction from the coherent amplitude analysis.	
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$4.9\pm 0.6\pm 0.5$	LINK 07A FOCS 4-body fit, $\approx 5.7$ k evts
$\Gamma(\pi^+\pi^-2\pi^0)/\Gamma(K^-\pi^+)$	$\Gamma_{166}/\Gamma_{31}$
<u>VALUE (units <math>10^{-2}</math>)</u> <u>EVTS</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$25.8\pm 1.5\pm 1.8$ 2724 $\pm$ 166	RUBIN 06 CLEO $e^+e^-\text{ at } \psi(3770)$
$\Gamma(\eta\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{167}/\Gamma$
Unseen decay modes of the $\eta$ are included.	
<u>VALUE (units <math>10^{-4}</math>)</u> <u>EVTS</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •	
6.4 $\pm$ 1.0 $\pm$ 0.4 156 $\pm$ 24	ARTUSO 08 CLEO See MENDEZ 10
$\Gamma(\eta\pi^0)/\Gamma(K^-\pi^+)$	$\Gamma_{167}/\Gamma_{31}$
Unseen decay modes of the $\eta$ are included.	
<u>VALUE (units <math>10^{-2}</math>)</u> <u>EVTS</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •	
1.47 $\pm$ 0.34 $\pm$ 0.11 62 $\pm$ 14	RUBIN 06 CLEO See ARTUSO 08
$\Gamma(\eta\pi^0)/[\Gamma(K^-\pi^+)+\Gamma(K^+\pi^-)]$	$\Gamma_{167}/(\Gamma_{31}+\Gamma_{222})$
Unseen decay modes of the $\eta$ are included.	
<u>VALUE (units <math>10^{-2}</math>)</u> <u>EVTS</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$1.74\pm 0.19$ OUR FIT	
$1.74\pm 0.15\pm 0.11$ 481 $\pm$ 40	MENDEZ 10 CLEO $e^+e^-\text{ at } 3774$ MeV

See key on page 601

Meson Particle Listings  
 $D^0$ 

$\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{168}/\Gamma$
Unseen decay modes of the $\omega$ are included.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.6 \times 10^{-4}$	90	RUBIN	06	CLEO $e^+e^-$ at $\psi(3770)$	

$\Gamma(2\pi^+2\pi^-\pi^0)/\Gamma(K^-\pi^+)$					$\Gamma_{169}/\Gamma_{31}$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
$10.7 \pm 1.2 \pm 0.5$	$1614 \pm 171$	RUBIN	06	CLEO $e^+e^-$ at $\psi(3770)$		

$\Gamma(\eta\pi^+\pi^-)/\Gamma_{\text{total}}$						$\Gamma_{170}/\Gamma$
Unseen decay modes of the $\eta$ are included.						
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>10.9 \pm 1.3 \pm 0.9</math></b>		$257 \pm 32$	ARTUSO	08	CLEO	$e^+e^-$ at $\psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$<19$	90		RUBIN	06	CLEO	$e^+e^-$ at $\psi(3770)$

$\Gamma(\omega\pi^+\pi^-)/\Gamma(K^-\pi^+)$					$\Gamma_{171}/\Gamma_{31}$
Unseen decay modes of the $\omega$ are included.					
$\frac{VALUE\ (units\ 10^{-2})}{4.1 \pm 1.2 \pm 0.4}$	$\frac{EVTS}{472 \pm 132}$	$\frac{DOCUMENT\ ID}{RUBIN}$	$\frac{TECN}{06}$	$\frac{COMMENT}{CLEO\ e^+e^- \text{ at } \psi(3770)}$	

$\Gamma(3\pi^+3\pi^-)/\Gamma(K^-\pi^+\pi^-)$					$\Gamma_{172}/\Gamma_{67}$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
$5.23 \pm 0.59 \pm 1.35$	$149 \pm 17$	LINK	04B	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV		

$\Gamma(3\pi^+3\pi^-)/\Gamma(K^-\pi^+\pi^-)$				$\Gamma_{172}/\Gamma_{93}$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.93 \pm 0.47 \pm 0.48$	<sup>1</sup> LINK	04B	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV	
<sup>1</sup> This LINK 04B result is not independent of other results in these Listings.				

$\Gamma(\eta'(958)\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{173}/\Gamma$	
Unseen decay modes of the $\eta'(958)$ are included.						
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$8.1 \pm 1.5 \pm 0.6$	$50 \pm 9$	ARTUSO	08	CLEO	See MENDEZ 10	

$\Gamma(\eta'(958)\pi^0)/[\Gamma(K^-\pi^+)+\Gamma(K^+\pi^-)]$					$\Gamma_{173}/(\Gamma_{31}+\Gamma_{222})$	
Unseen decay modes of the $\eta'(958)$ are included.						
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>2.3±0.4 OUR FIT</b>						
<b>2.3±0.3±0.2</b>	159 ± 19	MENDEZ	10	CLEO	$e^+e^-$ at 3774 MeV	

$\Gamma(\eta'(958)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{174}/\Gamma$
Unseen decay modes of the $\eta'(958)$ are included.					
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
$4.5 \pm 1.6 \pm 0.5$	$21 \pm 8$	ARTUSO	08	CLEO	$e^+e^-$ at $\psi(3770)$

$\Gamma(2\eta)/\Gamma_{\text{total}}$					$\Gamma_{175}/\Gamma$	
Unseen decay modes of the $\eta$ are included.						
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$16.7 \pm 1.4 \pm 1.3$	$255 \pm 22$	ARTUSO	08	CLEO	See MENDEZ 10	

$\Gamma(2\eta)/[\Gamma(K^-\pi^+) + \Gamma(K^+\pi^-)]$					$\Gamma_{175}/(\Gamma_{31}+\Gamma_{222})$	
Unseen decay modes of the $\eta$ are included.						
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>4.3±0.5 OUR FIT</b>						
<b>4.3±0.3±0.4</b>	430 ± 29	MENDEZ	10	CLEO	$e^+e^-$ at 3774 MeV	

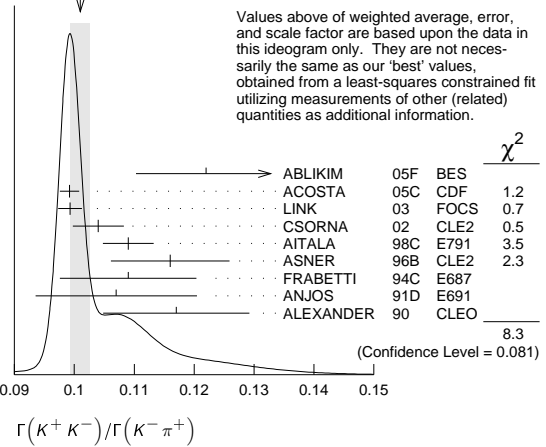
$\Gamma(\eta\eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{176}/\Gamma$
Unseen decay modes of the $\eta$ and $\eta'(958)$ are included.					
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$12.6 \pm 2.5 \pm 1.1$	$46 \pm 9$	ARTUSO	08	CLEO	See MENDEZ 10

$\Gamma(\eta\eta'(958))/[\Gamma(K^-\pi^+)+\Gamma(K^+\pi^-)]$				$\Gamma_{176}/(\Gamma_{31}+\Gamma_{222})$	
Unseen decay modes of the $\eta$ and $\eta'(958)$ are included.					
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.7±0.7 OUR FIT</b>					
<b>2.7±0.6±0.3</b>	66 ± 15	MENDEZ	10	CLEO	$e^+e^-$ at 3774 MeV

Hadronic modes with a  $K\bar{K}$  pair

$\Gamma(K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{177}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>4.01 ± 0.07 OUR FIT</b>	Error includes scale factor of 1.5.					
• • • We do not use the following data for averages, fits, limits, etc. • • •						
4.08 ± 0.08 ± 0.09	4746 ± 74	BONVICINI	08	CLEO	See MENDEZ 10	

$\Gamma(K^+K^-)/\Gamma(K^-\pi^+)$				$\Gamma_{177}/\Gamma_{31}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1021 ± 0.0015 OUR FIT</b>	Error includes scale factor of 1.7.			
<b>0.1010 ± 0.0016 OUR AVERAGE</b>	Error includes scale factor of 1.4. See the ideogram below.			
0.122 ± 0.011 ± 0.004	242 ± 20	ABLIKIM	05F BES	$e^+e^- \approx \psi(3770)$
0.0992 ± 0.0011 ± 0.0012	16k ± 200	ACOSTA	05c CDF	$p\bar{p}, \sqrt{s}=1.96$ TeV
0.0993 ± 0.0014 ± 0.0014	11k	LINK	03 FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV
0.1040 ± 0.0033 ± 0.0027	1900	CSORNA	02 CLE2	$e^+e^- \approx \Upsilon(4S)$
0.109 ± 0.003 ± 0.003	3317	AITALA	98c E791	$\pi^-$ nucleus, 500 GeV
0.116 ± 0.007 ± 0.007	1102	ASNER	96B CLE2	$e^+e^- \approx \Upsilon(4S)$
0.109 ± 0.007 ± 0.009	581	FRABETTI	94c E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
0.107 ± 0.010 ± 0.009	193	ANJOS	91D E691	Photoproduction
0.117 ± 0.010 ± 0.007	249	ALEXANDER	90 CLEO	$e^+e^-$ 10.5–11 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.107 ± 0.029 ± 0.015	103	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
0.138 ± 0.027 ± 0.010	155	FRABETTI	92 E687	$\gamma$ Be
0.16 ± 0.05	34	ALVAREZ	91B NA14	Photoproduction
0.10 ± 0.02 ± 0.01	131	ALBRECHT	90C ARG	$e^+e^- \approx 10$ GeV
0.122 ± 0.018 ± 0.012	118	BALTRUSAIT...	85E MRK3	$e^+e^-$ 3.77 GeV
0.113 ± 0.030		ABRAMS	79D MRK2	$e^+e^-$ 3.77 GeV

WEIGHTED AVERAGE  
0.1010±0.0016 (Error scaled by 1.4)

$\Gamma(K^+K^-)/[\Gamma(K^-\pi^+) + \Gamma(K^+\pi^-)]$					$\Gamma_{177}/(\Gamma_{31} + \Gamma_{222})$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>10.18 ± 0.15 OUR FIT</b>				Error includes scale factor of 1.7.		
$10.41 \pm 0.11 \pm 0.12$	13.8k	MENDEZ	10	CLEO $e^+e^-$ at 3774 MeV		

$\Gamma(K^+K^-)/\Gamma(\pi^+\pi^-)$				$\Gamma_{177}/\Gamma_{132}$	
The unused results here are redundant with $\Gamma(K^+K^-)/\Gamma(K^-\pi^+)$ and $\Gamma(\pi^+\pi^-)/\Gamma(K^-\pi^+)$ measurements by the same experiments.					
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$2.760 \pm 0.040 \pm 0.034$	7334	ACOSTA	05c	CDF	$p\bar{p}, \sqrt{s}=1.96$ TeV
$2.81 \pm 0.10 \pm 0.06$		LINK	03	FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV
$2.96 \pm 0.16 \pm 0.15$	710	CSORNA	02	CLE2	$e^+e^- \approx \Upsilon(4S)$
$2.75 \pm 0.15 \pm 0.16$		AITALA	98c	E791	$\pi^-$ nucleus, 500 GeV
$2.53 \pm 0.46 \pm 0.19$		FRABETTI	94c	E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
$2.23 \pm 0.81 \pm 0.46$		ADAMOVICH	92	OMEG	$\pi^-$ 340 GeV
$1.95 \pm 0.34 \pm 0.22$		ANJOS	91D	E691	Photoproduction
$2.5 \pm 0.7$		ALBRECHT	90c	ARG	$e^+e^- \approx 10$ GeV
$2.35 \pm 0.37 \pm 0.28$		ALEXANDER	90	CLEO	$e^+e^-$ 10.5–11 GeV

$\Gamma(2K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{178}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$1.46 \pm 0.32 \pm 0.09$	$68 \pm 15$	BONVICINI	08	CLEO	See MENDEZ 10

$\Gamma(2K_S^0)/[\Gamma(K^-\pi^+) + \Gamma(K^+\pi^-)]$					$\Gamma_{178}/(\Gamma_{31} + \Gamma_{222})$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>0.45 ± 0.11 OUR FIT</b>				Error includes scale factor of 2.5.		
$0.41 \pm 0.04 \pm 0.02$	$215 \pm 23$	MENDEZ	10	CLEO $e^+e^-$ at 3774 MeV		

# Meson Particle Listings

$D^0$

$\Gamma(2K_S^0)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{178}/\Gamma_{35}$   
This is the same as  $\Gamma(K^0\bar{K}^0)/\Gamma(\bar{K}^0\pi^+\pi^-)$  because  $D^0\rightarrow K_S^0K_L^0$  is forbidden by  $CP$  conservation.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0062±0.0015 OUR FIT</b>	Error includes scale factor of 2.2.			
<b>0.0120±0.0022 OUR AVERAGE</b>				
0.0144±0.0032±0.0016	79 ± 17	LINK	05A FOCs	$\gamma$ Be, $\bar{E}_\gamma \approx 180$ GeV
0.0101±0.0022±0.0016	26	ASNER	96B CLE2	$e^+e^- \approx \gamma(4S)$
0.039 ± 0.013 ± 0.013	20 ± 7	FRABETTI	94J E687	$\gamma$ Be $\bar{E}_\gamma=220$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.021 <sup>+0.011</sup> <sub>-0.008</sub> ± 0.002	5	ALEXANDER	90 CLEO	$e^+e^-$ 10.5–11 GeV

$\Gamma(K_S^0K^-\pi^+)/\Gamma(K^-\pi^+)$   $\Gamma_{179}/\Gamma_{31}$   

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.091±0.014 OUR FIT</b>	Error includes scale factor of 1.2.		
<b>0.08 ± 0.03</b>	<sup>1</sup> ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>1</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K_S^0K^-\pi^+)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{179}/\Gamma_{35}$   

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.125±0.017 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.119±0.021 OUR AVERAGE</b>	Error includes scale factor of 1.3.			
0.108±0.019	61	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
0.16 ± 0.03 ± 0.02	39	ALBRECHT	90c ARG	$e^+e^- \approx 10$ GeV

$\Gamma(\bar{K}^*(892)^0K_S^0, \bar{K}^{*0} \rightarrow K^-\pi^+)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{180}/\Gamma_{35}$   

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.019</b>	90	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.02	90	ALBRECHT	90c ARG	$e^+e^- \approx 10$ GeV

$\Gamma(K_S^0K^+\pi^-)/\Gamma(K^-\pi^+)$   $\Gamma_{181}/\Gamma_{31}$   

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.055±0.009 OUR FIT</b>	Error includes scale factor of 1.3.		
<b>0.05 ± 0.025</b>	<sup>1</sup> ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>1</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K_S^0K^+\pi^-)/\Gamma(K_S^0\pi^+\pi^-)$   $\Gamma_{181}/\Gamma_{35}$   

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.076±0.012 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>0.098±0.020</b>	55	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV

$\Gamma(K_S^0K^+\pi^-)/\Gamma(K_S^0K^-\pi^+)$   $\Gamma_{181}/\Gamma_{179}$   

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.61 ± 0.06 OUR FIT</b>	Error includes scale factor of 1.3.		
<b>0.592±0.044±0.018</b>	INSLER	12 CLEO	$e^+e^- \rightarrow D^0\bar{D}^0$ at 3.77 GeV

$\Gamma(K^*(892)^0K_S^0, K^{*0} \rightarrow K^+\pi^-)/\Gamma(\bar{K}^*(892)^0K_S^0, \bar{K}^{*0} \rightarrow K^-\pi^+)$   $\Gamma_{182}/\Gamma_{180}$   

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.356±0.034±0.007</b>	<sup>1</sup> INSLER	12 CLEO	$e^+e^- \rightarrow D^0\bar{D}^0$ , 3.77 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.010	90	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV

<sup>1</sup> Uses quantum correlations in  $e^+e^- \rightarrow D^0\bar{D}^0$  at the  $\psi(3770)$ , where the signal side  $D$  decays to  $K_S^0K\pi$  and the tag-side  $D$  decays to  $K\pi, K\pi\pi\pi, K\pi\pi^0$ .

$\Gamma(K^+K^-\pi^0)/\Gamma(K^-\pi^+\pi^0)$   $\Gamma_{183}/\Gamma_{50}$   

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.37±0.03±0.04</b>	11k±122	AUBERT,B	06x BABR	$e^+e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.95±0.26	151	ASNER	96B CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(K^*(892)^+K^-, K^*(892)^+ \rightarrow K^+\pi^0)/\Gamma(K^+K^-\pi^0)$   $\Gamma_{184}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>44.4±0.8±0.6</b>	AUBERT	07T BABR	Dalitz fit II, 11k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
46.1±3.1	<sup>1</sup> CAWLFIELD	06A CLEO	Dalitz fit, 627 ± 30 evts

<sup>1</sup> The error on this CAWLFIELD 06A result is statistical only.

$\Gamma(K^*(892)^-K^+, K^*(892)^- \rightarrow K^-\pi^0)/\Gamma(K^+K^-\pi^0)$   $\Gamma_{185}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>15.9±0.7±0.6</b>	AUBERT	07T BABR	Dalitz fit II, 11k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
12.3±2.2	<sup>1</sup> CAWLFIELD	06A CLEO	Dalitz fit, 627 ± 30 evts

<sup>1</sup> The error on this CAWLFIELD 06A result is statistical only.

$\Gamma((K^+\pi^0)S\text{-wave}K^-)/\Gamma(K^+K^-\pi^0)$   $\Gamma_{186}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>71.1±3.7±1.9</b>	<sup>1</sup> AUBERT	07T BABR	Dalitz fit II, 11k evts

<sup>1</sup> The only major difference between fits I and II in the AUBERT 07T analysis is in this mode, where the fit-I fraction is (16.3 ± 3.4 ± 2.1)%.

$\Gamma((K^-\pi^0)S\text{-wave}K^+)/\Gamma(K^+K^-\pi^0)$   $\Gamma_{187}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>3.9±0.9±1.0</b>	AUBERT	07T BABR	Dalitz fit II, 11k evts

$\Gamma(f_0(980)\pi^0, f_0 \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^0)$   $\Gamma_{188}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>10.5±1.1±1.2</b>	<sup>1</sup> AUBERT	07T BABR	Dalitz fit II, 11k evts

<sup>1</sup> When AUBERT 07T replace the  $f_0(980)\pi^0$  mode with  $a_0(980)\pi^0$ , the fit fraction is a negligibly different (11.0 ± 1.5 ± 1.2)%.

$\Gamma(\phi\pi^0, \phi \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^0)$   $\Gamma_{189}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>19.4±0.6±0.5</b>	AUBERT	07T BABR	Dalitz fit II, 11k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
14.9±1.6	<sup>1</sup> CAWLFIELD	06A CLEO	Dalitz fit, 627 ± 30 evts

<sup>1</sup> The error on this CAWLFIELD 06A result is statistical only.

$\Gamma(K^+K^-\pi^0\text{ nonresonant})/\Gamma(K^+K^-\pi^0)$   $\Gamma_{190}/\Gamma_{183}$   
This is the “fit fraction” from the Dalitz-plot analysis with interference.  

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.360±0.037	<sup>1</sup> CAWLFIELD	06A CLEO	Dalitz fit, 627 ± 30 evts

<sup>1</sup> The error is statistical only. CAWLFIELD 06A also fits the Dalitz plot replacing this flat nonresonant background with broad  $S\text{-wave } \kappa^\pm \rightarrow K^\pm\pi^0$  resonances. There is no significant improvement in the fit, and  $K^*\pm K^\mp$  and  $\phi\pi^0$  results are not much changed.

$\Gamma(2K_S^0\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{191}/\Gamma$   

VALUE	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.00059</b>	ASNER	96B CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(\phi\pi^0)/\Gamma(K^+K^-)$   $\Gamma_{213}/\Gamma_{177}$   

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.194±0.006±0.009	1254	TAJIMA	04 BELL	$e^+e^-$ at $\gamma(4S)$

$\Gamma(\phi\eta)/\Gamma(K^+K^-)$   $\Gamma_{214}/\Gamma_{177}$   

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.59±1.14±0.18</b>	31	TAJIMA	04 BELL	$e^+e^-$ at $\gamma(4S)$

$\Gamma(\phi\omega)/\Gamma_{\text{total}}$   $\Gamma_{215}/\Gamma$   

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.0021</b>	90	ALBRECHT	94i ARG	$e^+e^- \approx 10$ GeV

$\Gamma(K^+K^-\pi^+\pi^-)/\Gamma(K^-2\pi^+\pi^-)$   $\Gamma_{192}/\Gamma_{67}$   

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.00±0.13 OUR AVERAGE</b>				
2.95±0.11±0.08	2669 ± 101	<sup>1</sup> LINK	05G FOCs	$\gamma$ Be, $\bar{E}_\gamma \approx 180$ GeV
3.13±0.37±0.36	136 ± 15	AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
3.5 ± 0.4 ± 0.2	244 ± 26	FRABETTI	95c E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.4 ± 1.8 ± 0.5	19 ± 8	ABLIKIM	05F BES	$e^+e^- \approx \psi(3770)$
4.1 ± 0.7 ± 0.5	114 ± 20	ALBRECHT	94i ARG	$e^+e^- \approx 10$ GeV
3.14±1.0	89 ± 29	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
2.8 <sup>+0.8</sup> <sub>-0.7</sub>		ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>1</sup> LINK 05G uses a smaller, cleaner subset of 1279 ± 48 events for the amplitude analysis that gives the results in the next data blocks.

$\Gamma(\phi(\pi^+\pi^-)S\text{-wave}, \phi \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^+\pi^-)$   $\Gamma_{193}/\Gamma_{192}$   
This is the fraction from a coherent amplitude analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>10.3±1.0±0.8</b>	ARTUSO	12 CLEO	Fitting 2959 evts.
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1 ± 1	LINK	05G FOCs	Fits 1279 ± 48 evts.

$\Gamma((\phi\rho^0)S\text{-wave}, \phi \rightarrow K^+K^-)/\Gamma(K^+K^-\pi^+\pi^-)$   $\Gamma_{194}/\Gamma_{192}$   
This is the fraction from a coherent amplitude analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>38.3±2.5±3.8</b>	ARTUSO	12 CLEO	Fitting 2959 evts.
• • • We do not use the following data for averages, fits, limits, etc. • • •			
29 ± 2 ± 1	LINK	05G FOCs	Fits 1279 ± 48 evts.

$\Gamma((\phi\rho^0)_{D\text{-wave}}, \phi \rightarrow K^+ K^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{195}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$3.4 \pm 0.7 \pm 0.6$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma((K^{*0}\bar{K}^{*0})_{S\text{-wave}}, K^{*0} \rightarrow K^\pm \pi^\mp) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{196}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$6.1 \pm 0.8 \pm 0.9$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma((K^- \pi^+)_{P\text{-wave}}, (K^+ \pi^-)_{S\text{-wave}}) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{197}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$10.9 \pm 1.2 \pm 1.7$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K_1(1270)^+ K^-, K_1(1270)^+ \rightarrow K^{*0} \pi^+) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{198}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$7.3 \pm 0.8 \pm 1.9$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K_1(1270)^+ K^-, K_1(1270)^+ \rightarrow \rho^0 K^+) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{199}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$4.7 \pm 0.7 \pm 0.8$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K_1(1270)^- K^+, K_1(1270)^- \rightarrow \bar{K}^{*0} \pi^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{200}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$0.9 \pm 0.3 \pm 0.4$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K_1(1270)^- K^+, K_1(1270)^- \rightarrow \rho^0 K^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{201}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$6.0 \pm 0.8 \pm 0.6$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K^*(1410)^+ K^-, K^*(1410)^+ \rightarrow K^{*0} \pi^+) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{202}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$4.2 \pm 0.7 \pm 0.8$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K^*(1410)^- K^+, K^*(1410)^- \rightarrow \bar{K}^{*0} \pi^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{203}/\Gamma_{192}$
VALUE (%)	DOCUMENT ID
$4.7 \pm 0.7 \pm 0.7$	ARTUSO 12 CLEO Fitting 2959 evts.

$\Gamma(K^+ K^- \rho^0 \text{ 3-body}) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{204}/\Gamma_{192}$
This is the fraction from a coherent amplitude analysis.	
VALUE (%)	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$2 \pm 2 \pm 2$	LINK 05G FOCS Fits 1279 $\pm$ 48 evts.

$\Gamma(f_0(980) \pi^\pm \pi^-, f_0 \rightarrow K^+ K^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{205}/\Gamma_{192}$
This is the fraction from a coherent amplitude analysis.	
VALUE (%)	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$15 \pm 3 \pm 2$	LINK 05G FOCS Fits 1279 $\pm$ 48 evts.

$\Gamma(K^*(892)^0 K^\mp \pi^\pm \text{ 3-body}, K^{*0} \rightarrow K^\pm \pi^\mp) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{206}/\Gamma_{192}$
This is the fraction from a coherent amplitude analysis.	
VALUE (%)	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$11 \pm 2 \pm 1$	LINK 05G FOCS Fits 1279 $\pm$ 48 evts.

$\Gamma(K^*(892)^0 \bar{K}^*(892)^0, K^{*0} \rightarrow K^\pm \pi^\mp) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{207}/\Gamma_{192}$
This is the fraction from a coherent amplitude analysis.	
VALUE (%)	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$3 \pm 2 \pm 1$	LINK 05G FOCS Fits 1279 $\pm$ 48 evts.

$\Gamma(K_1(1270)^\pm K^\mp, K_1(1270)^\pm \rightarrow K^\pm \pi^\mp \pi^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{208}/\Gamma_{192}$
This is the fraction from a coherent amplitude analysis.	
VALUE (%)	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$33 \pm 6 \pm 4$	<sup>1</sup> LINK 05G FOCS Fits 1279 $\pm$ 48 evts.

<sup>1</sup> This LINK 05G value includes  $K_1(1270)^\pm \rightarrow \rho^0 K^\pm \rightarrow K_0^*(1430)^0 \pi^\pm$ , and  $K^*(892)^0 \pi^\pm$ .

$\Gamma(K_1(1400)^\pm K^\mp, K_1(1400)^\pm \rightarrow K^\pm \pi^\mp \pi^-) / \Gamma(K^+ K^- \pi^+ \pi^-)$	$\Gamma_{209}/\Gamma_{192}$
This is the fraction from a coherent amplitude analysis.	
VALUE (%)	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$22 \pm 3 \pm 4$	LINK 05G FOCS Fits 1279 $\pm$ 48 evts.

$\Gamma(2K_S^0 \pi^+ \pi^-) / \Gamma(K_S^0 \pi^+ \pi^-)$	$\Gamma_{210}/\Gamma_{35}$
VALUE (units $10^{-2}$ )	DOCUMENT ID
$4.3 \pm 0.8$ OUR AVERAGE	
$4.16 \pm 0.70 \pm 0.42$	113 $\pm$ 21 LINK 05A FOCS $\gamma$ Be, $\bar{E}_\gamma \approx 180$ GeV
$6.2 \pm 2.0 \pm 1.6$	25 ALBRECHT 94i ARG $e^+ e^- \approx 10$ GeV

$\Gamma(K_S^0 K^- 2\pi^+ \pi^-) / \Gamma(K_S^0 2\pi^+ 2\pi^-)$	$\Gamma_{211}/\Gamma_{87}$
VALUE	DOCUMENT ID
$<0.054$	90 LINK 04D FOCS $\gamma$ A, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) / \Gamma_{\text{total}}$	$\Gamma_{212}/\Gamma$
VALUE	DOCUMENT ID
$0.0031 \pm 0.0020$	<sup>1</sup> BARLAG 92c ACCM $\pi^-$ Cu 230 GeV
<sup>1</sup> BARLAG 92c computes the branching fraction using topological normalization.	

## Radiative modes

$\Gamma(\rho^0 \gamma) / \Gamma_{\text{total}}$	$\Gamma_{216}/\Gamma$
VALUE	DOCUMENT ID
$<2.4 \times 10^{-4}$	90 ASNER 98 CLE2

$\Gamma(\omega \gamma) / \Gamma_{\text{total}}$	$\Gamma_{217}/\Gamma$
VALUE	DOCUMENT ID
$<2.4 \times 10^{-4}$	90 ASNER 98 CLE2

$\Gamma(\phi \gamma) / \Gamma(K^+ K^-)$	$\Gamma_{218}/\Gamma_{177}$
VALUE (units $10^{-3}$ )	DOCUMENT ID
$6.8 \pm 0.9$ OUR FIT	
$6.31 \pm 1.70 \pm 0.30$	28 TAJIMA 04 BELL $e^+ e^-$ at $\gamma(4S)$
$-1.48 - 0.36$	

$\Gamma(\phi \gamma) / \Gamma(K^- \pi^+)$	$\Gamma_{218}/\Gamma_{31}$
VALUE (units $10^{-4}$ )	DOCUMENT ID
$7.0 \pm 0.9$ OUR FIT	
$7.15 \pm 0.78 \pm 0.69$	243 $\pm$ 25 AUBERT 08AZ BABR $e^+ e^- \approx 10.6$ GeV

$\Gamma(\bar{K}^*(892)^0 \gamma) / \Gamma(K^- \pi^+)$	$\Gamma_{219}/\Gamma_{31}$
VALUE (units $10^{-3}$ )	DOCUMENT ID
$8.43 \pm 0.51 \pm 0.70$	2286 $\pm$ 113 AUBERT 08AZ BABR $e^+ e^- \approx 10.6$ GeV

## Doubly Cabibbo-suppressed / Mixing modes

$\Gamma(K^+ \ell^- \bar{\nu}_\ell \text{ via } \bar{D}^0) / \Gamma(K^- \ell^+ \nu_\ell)$	$\Gamma_{220}/\Gamma_{17}$
This is a limit on $R_M$ without the complications of possible doubly Cabibbo-suppressed decays that occur when using hadronic modes. For the limits on $ m_1 - m_2 $ and $(\Gamma_1 - \Gamma_2)/\Gamma$ that come from the best mixing limit, see near the beginning of these $D^0$ Listings.	

VALUE	DOCUMENT ID
$< 6.1 \times 10^{-4}$	<sup>1</sup> BITENC 08 BELL $e^+ e^-$ , 10.58 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<50 \times 10^{-4}$	<sup>2</sup> AITALA 96c E791 $\pi^-$ nucleus, 500 GeV

<sup>1</sup> The BITENC 08 right-sign sample includes about 15% of  $D^0 \rightarrow K^- \pi^0 \ell^+ \nu_\ell$  and other decays.

<sup>2</sup> AITALA 96c uses  $D^{*+} \rightarrow D^0 \pi^+$  (and charge conjugate) decays to identify the charm at production and  $D^0 \rightarrow K^- \ell^+ \nu_\ell$  (and charge conjugate) decays to identify the charm at decay.

$\Gamma(K^+ \alpha K^*(892)^+ e^- \bar{\nu}_e \text{ via } \bar{D}^0) / [\Gamma(K^- e^+ \nu_e) + \Gamma(K^*(892)^- e^+ \nu_e)]$	$\Gamma_{221}/(\Gamma_{18} + \Gamma_{20})$
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This is a limit on  $R_M$  without the complications of possible doubly Cabibbo-suppressed decays that occur when using hadronic modes. The experiments use  $D^{*+} \rightarrow D^0 \pi^+$  (and charge conjugate) decays to identify the charm at production and the charge of the  $e$  to identify the charm at decay. These limits do not allow  $CP$  violation. For the limits on  $|m_1 - m_2|$  and  $(\Gamma_1 - \Gamma_2)/\Gamma$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	DOCUMENT ID
$<0.001$	90 BITENC 05 BELL $e^+ e^- \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$-0.0013 < R < +0.0012$	90 AUBERT 07AB BABR $e^+ e^- \approx 10.58$ GeV
$<0.0078$	90 CAWLFIELD 05 CLEO $e^+ e^- \approx 10.6$ GeV
$<0.0042$	90 AUBERT,B 04Q BABR See AUBERT 07AB

$\Gamma(K^+ \pi^-) / \Gamma(K^- \pi^+)$	$\Gamma_{222}/\Gamma_{31}$
This is $R$ , the time-integrated wrong-sign rate compared to the right-sign rate. See the note on " $D^0$ - $\bar{D}^0$ Mixing," near the start of the $D^0$ Listings.	

The experiments here use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. The  $D^0 \rightarrow K^+ \pi^-$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+ \pi^-$  decay. Some of the experiments can use the decay-time information to disentangle the two mechanisms. Here, we list the experimental branching ratio, which if there is no mixing is the DCS ratio. See the next data block for values of the DCS ratio  $R_D$ , and the following data block for limits on the mixing ratio  $R_M$ . See the section on  $CP$ -violating asymmetries near the end of this  $D^0$  Listing for values of  $A_D$ , and the note on " $D^0$ - $\bar{D}^0$  Mixing" for limits on  $x'$  and  $y'$ .

Some early limits have been omitted from this Listing; see our 1998 edition (The European Physical Journal **C3** (1998)) and our 2006 edition (Journal of Physics **G33** 1 (2006)).

VALUE (units $10^{-3}$ )	DOCUMENT ID
$3.79 \pm 0.18$ OUR FIT	Error includes scale factor of 3.3.
$3.79 \pm 0.18$ OUR AVERAGE	Error includes scale factor of 3.3. See the ideogram below.
$4.15 \pm 0.10$	<sup>1</sup> AALTONEN 08E CDF $p\bar{p}$ , $\sqrt{s} = 1.96$ TeV
$3.53 \pm 0.08 \pm 0.04$	<sup>2</sup> AUBERT 07W BABR $e^+ e^- \approx 10.6$ GeV
$3.77 \pm 0.08 \pm 0.05$	<sup>1</sup> ZHANG 06 BELL $e^+ e^-$

## Meson Particle Listings

 $D^0$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.05 \pm 0.21 \pm 0.11$	$2.0 \pm 0.1k$	<sup>3</sup> ABULENCIA	06x	CDF	See AALTONEN 08E
$3.81 \pm 0.17 \pm 0.08$ $-0.16$	$845 \pm 40$	<sup>2</sup> LI	05A	BELL	See ZHANG 06
$4.29 \pm 0.63 \pm 0.27$	234	<sup>4</sup> LINK	05H	FOCS	$\gamma$ nucleus
$3.57 \pm 0.22 \pm 0.27$		<sup>5</sup> AUBERT	03Z	BABR	See AUBERT 07W
$4.04 \pm 0.85 \pm 0.25$	149	<sup>6</sup> LINK	01	FOCS	$\gamma$ nucleus
$3.32 \pm 0.63 \pm 0.40$ $-0.65$	45	<sup>1</sup> GODANG	00	CLE2	$e^+e^-$
$6.8 \pm 3.4 \pm 0.7$ $-3.3$	34	<sup>2</sup> AITALA	98	E791	$\pi^-$ nucl., 500 GeV

<sup>1</sup> GODANG 00, ZHANG 06, and AALTONEN 08E allow  $CP$  violation.

<sup>2</sup> AITALA 98, LI 05A, and AUBERT 07W assume no  $CP$  violation.

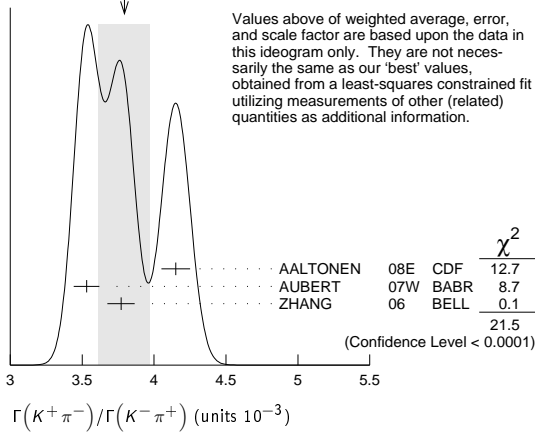
<sup>3</sup> This ABULENCIA 06x result assumes no mixing.

<sup>4</sup> This LINK 05H result assumes no mixing but allows  $CP$  violation. If neither mixing nor  $CP$  violation is allowed,  $R = (4.29 \pm 0.63 \pm 0.28) \times 10^{-3}$ .

<sup>5</sup> This AUBERT 03Z result allows  $CP$  violation. If  $CP$  violation is not allowed,  $R = 0.00359 \pm 0.00020 \pm 0.00027$ .

<sup>6</sup> This LINK 01 result assumes no mixing or  $CP$  violation.

WEIGHTED AVERAGE  
 $3.79 \pm 0.18$  (Error scaled by 3.3)



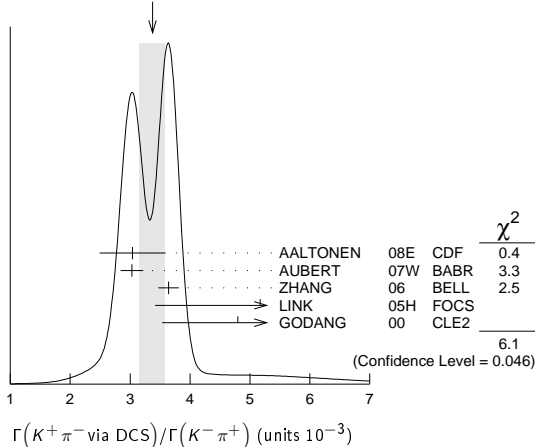
$\Gamma(K^+\pi^- \text{ via DCS})/\Gamma(K^-\pi^+)$

$\Gamma_{223}/\Gamma_{31}$

This is  $R_D$ , the doubly Cabibbo-suppressed ratio when mixing is allowed.

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.37 \pm 0.21</math> OUR AVERAGE</b>		Error includes scale factor of 1.8. See the ideogram below.			
$3.04 \pm 0.55$		$12.7 \pm 0.3k$	AALTONEN	08E	CDF $p\bar{p}$ , $\sqrt{s} = 1.96$ TeV
$3.03 \pm 0.16 \pm 0.10$		$4030 \pm 90$	<sup>1</sup> AUBERT	07W	BABR $e^+e^- \approx 10.6$ GeV
$3.64 \pm 0.17$		$4024 \pm 88$	<sup>2</sup> ZHANG	06	BELL $e^+e^-$
$5.17 \pm 1.47 \pm 0.76$ $-1.58$		234	<sup>3</sup> LINK	05H	FOCS $\gamma$ nucleus
$4.8 \pm 1.2 \pm 0.4$		45	<sup>4</sup> GODANG	00	CLE2 $e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$2.87 \pm 0.37$		$845 \pm 40$	LI	05A	BELL See ZHANG 06
$2.3 < R_D < 5.2$	95		<sup>5</sup> AUBERT	03Z	BABR See AUBERT 07W
$9.0 \pm 12.0 \pm 4.4$ $-10.9$		34	<sup>6</sup> AITALA	98	E791 $\pi^-$ nucl., 500 GeV

WEIGHTED AVERAGE  
 $3.37 \pm 0.21$  (Error scaled by 1.8)



$\Gamma(K^+\pi^- \text{ via DCS})/\Gamma(K^-\pi^+)$  (units  $10^{-3}$ )

<sup>1</sup> This AUBERT 07W result is the same whether or not  $CP$  violation is allowed.

<sup>2</sup> This ZHANG 06 assumes no  $CP$  violation.

<sup>3</sup> This LINK 05H result allows  $CP$  violation. Allowing mixing but not  $CP$  violation,  $R_D = (3.81 \pm 1.67 \pm 0.92) \times 10^{-3}$ .

<sup>4</sup> This GODANG 00 result allows  $CP$  violation.

<sup>5</sup> This AUBERT 03Z result allows  $CP$  violation. If only mixing is allowed, the 95% confidence level interval is  $(2.4 < R_D < 4.9) \times 10^{-3}$ .

<sup>6</sup> This AITALA 98 result assumes no  $CP$  violation.

$\Gamma(K^+\pi^- \text{ via } \bar{D}^0)/\Gamma(K^-\pi^+)$

$\Gamma_{224}/\Gamma_{31}$

This is  $R_M$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings. The experiments here (1) use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_1 - m_2|$  and  $(\Gamma_1 - \Gamma_2)/\Gamma$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.00040</math></b>	95	<sup>1</sup> ZHANG	06	BELL $e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<0.00046$	95	<sup>2</sup> LI	05A	BELL See ZHANG 06
$<0.0063$	95	<sup>3</sup> LINK	05H	FOCS $\gamma$ nucleus
$<0.0013$	95	<sup>4</sup> AUBERT	03Z	BABR $e^+e^-$ , 10.6 GeV
$<0.00041$	95	<sup>5</sup> GODANG	00	CLE2 $e^+e^-$
$<0.0092$	95	<sup>6</sup> BARATE	98W	ALEP $e^+e^-$ at $Z^0$
$<0.005$	90	<sup>7</sup> ANJOS	88c	E691 Photoproduction

<sup>1</sup> This ZHANG 06 result allows  $CP$  violation, but the result does not change if  $CP$  violation is not allowed.

<sup>2</sup> This LI 05A result allows  $CP$  violation. The limit becomes  $< 0.00042$  (95% CL) if  $CP$  violation is not allowed.

<sup>3</sup> LINK 05H obtains the same result whether or not  $CP$  violation is allowed.

<sup>4</sup> This AUBERT 03Z result allows  $CP$  violation and assumes that the strong phase between  $D^0 \rightarrow K^+\pi^-$  and  $\bar{D}^0 \rightarrow K^+\pi^-$  is small, and limits only  $D^0 \rightarrow \bar{D}^0$  transitions via off-shell intermediate states. The limit on transitions via on-shell intermediate states is 0.0016.

<sup>5</sup> This GODANG 00 result allows  $CP$  violation and assumes that the strong phase between  $D^0 \rightarrow K^+\pi^-$  and  $\bar{D}^0 \rightarrow K^+\pi^-$  is small, and limits only  $D^0 \rightarrow \bar{D}^0$  transitions via off-shell intermediate states. The limit on transitions via on-shell intermediate states is 0.0017.

<sup>6</sup> This BARATE 98W result assumes no interference between the DCS and mixing amplitudes ( $y' = 0$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.036 (95% CL).

<sup>7</sup> This ANJOS 88c result assumes no interference between the DCS and mixing amplitudes ( $y' = 0$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.019.

$\Gamma(K_S^0 \pi^+ \pi^- \text{ in } D^0 \rightarrow \bar{D}^0)/\Gamma(K_S^0 \pi^+ \pi^-)$

$\Gamma_{225}/\Gamma_{35}$

This is  $R_M$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings. The experiments here (1) use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_1 - m_2|$  and  $(\Gamma_1 - \Gamma_2)/\Gamma$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.0063</math></b>	95	<sup>1</sup> ASNER	05	CLEO $e^+e^- \approx 10$ GeV

<sup>1</sup> This ASNER 05 limit allows  $CP$  violation. If  $CP$  violation is not allowed, the limit is 0.0042 at 95% CL.

$\Gamma(K^+\pi^-\pi^0)/\Gamma(K^-\pi^+\pi^0)$

$\Gamma_{229}/\Gamma_{50}$

The experiments here use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. The  $D^0 \rightarrow K^+\pi^-\pi^0$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+\pi^-\pi^0$  decay.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.20 \pm 0.10</math> OUR AVERAGE</b>				
$2.14 \pm 0.08 \pm 0.08$	$763 \pm 51$	<sup>1</sup> AUBERT,B	06N	BABR $e^+e^- \approx 7(45)$
$2.29 \pm 0.15 \pm 0.13$ $-0.09$	$1978 \pm 104$	TIAN	05	BELL $e^+e^- \approx 7(45)$
$4.3 \pm 1.1 \pm 0.7$ $-1.0$	38	BRANDENB...	01	CLE2 $e^+e^- \approx 7(45)$

<sup>1</sup> This AUBERT,B 06N result assumes no mixing.

$\Gamma(K^+\pi^-\pi^0 \text{ via } \bar{D}^0)/\Gamma(K^-\pi^+\pi^0)$

$\Gamma_{230}/\Gamma_{50}$

This is  $R_M$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings. The experiments here (1) use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_1 - m_2|$  and  $(\Gamma_1 - \Gamma_2)/\Gamma$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.25 \pm 0.25 \pm 0.12</math></b>		AUBERT	09AN	BABR $e^+e^-$ at 10.58 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.54$	95	<sup>1</sup> AUBERT,B	06N	BABR $e^+e^- \approx 7(45)$
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<sup>1</sup> This AUBERT,B 06N limit assumes no  $CP$  violation. The measured value corresponding to the limit is  $(2.3 \pm 1.8 \pm 0.4) \times 10^{-4}$ . If  $CP$  violation is allowed, this becomes  $(1.0 \pm 2.2 \pm 0.3) \times 10^{-4}$ .

$\Gamma(K^+\pi^+2\pi^-)/\Gamma(K^-2\pi^+\pi^-)$  $\Gamma_{231}/\Gamma_{67}$ 

The experiments here use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. The  $D^0 \rightarrow K^+\pi^-\pi^+\pi^-$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+\pi^-\pi^+\pi^-$  decay. Some of the experiments can use the decay-time information to disentangle the two mechanisms. Here, we list the experimental branching ratio, which if there is no mixing is the DCS ratio; in the next data block we give the limits on the mixing ratio.

Some early limits have been omitted from this Listing; see our 1998 edition (EPJ C3 1).

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.25 \pm 0.11</math></b>	<b>OUR AVERAGE</b>				
$3.24 \pm 0.08 \pm 0.07$		$3358 \pm 79$	<sup>1</sup> WHITE	13 BELL	$e^+e^- \approx \mathcal{T}(4S)$
$4.4^{+1.3}_{-1.2} \pm 0.4$		54	<sup>1</sup> DYTMAN	01 CLE2	$e^+e^- \approx \mathcal{T}(4S)$
$2.5^{+3.6}_{-3.4} \pm 0.3$			<sup>2</sup> AITALA	98 E791	$\pi^-$ nucl., 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.20 \pm 0.18^{+0.18}_{-0.13}$		$1721 \pm 75$	<sup>1</sup> TIAN	05 BELL	See WHITE 13
<18	90		<sup>1</sup> AMMAR	91 CLEO	$e^+e^- \approx 10.5$
<18	90		<sup>3</sup> ANJOS	88c E691	Photoproduction

<sup>1</sup> AMMAR 91 cannot and DYTMAN 01, TIAN 05, and WHITE 13 do not distinguish between doubly Cabibbo-suppressed decay and  $D^0\text{-}\bar{D}^0$  mixing.

<sup>2</sup> This AITALA 98 result assumes no  $D^0\text{-}\bar{D}^0$  mixing ( $R_M$  in the note on “ $D^0\text{-}\bar{D}^0$  Mixing”). It becomes  $-0.0020^{+0.0117}_{-0.0106} \pm 0.0035$  when mixing is allowed and decay-time information is used to distinguish doubly Cabibbo-suppressed decays from mixing.

<sup>3</sup> ANJOS 88c uses decay-time information to distinguish doubly Cabibbo-suppressed (DCS) decays from  $D^0\text{-}\bar{D}^0$  mixing. However, the result assumes no interference between the DCS and mixing amplitudes ( $\gamma = 0$  in the note on “ $D^0\text{-}\bar{D}^0$  Mixing” near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.033.

 $\Gamma(K^+\pi^+2\pi^- \text{ via } \bar{D}^0)/\Gamma(K^-2\pi^+\pi^-)$  $\Gamma_{232}/\Gamma_{67}$ 

This is a  $D^0\text{-}\bar{D}^0$  mixing limit. The experiments here (1) use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma_{D^0}$  that come from

the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.005</b>	90	<sup>1</sup> ANJOS	88c E691	Photoproduction

<sup>1</sup> ANJOS 88c uses decay-time information to distinguish doubly Cabibbo-suppressed (DCS) decays from  $D^0\text{-}\bar{D}^0$  mixing. However, the result assumes no interference between the DCS and mixing amplitudes ( $\gamma = 0$  in the note on “ $D^0\text{-}\bar{D}^0$  Mixing” near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.007.

 $\Gamma(K^+\pi^- \text{ or } K^+\pi^+2\pi^- \text{ via } \bar{D}^0)/\Gamma(K^-\pi^+ \text{ or } K^-2\pi^+\pi^-)$  $\Gamma_{233}/\Gamma_0$ 

This is a  $D^0\text{-}\bar{D}^0$  mixing limit. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma_{D^0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.0085	90	<sup>1</sup> AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
<0.0037	90	<sup>2</sup> ANJOS	88c E691	Photoproduction

<sup>1</sup> AITALA 98 uses decay-time information to distinguish doubly Cabibbo-suppressed decays from  $D^0\text{-}\bar{D}^0$  mixing. The fit allows interference between the two amplitudes, and also allows  $CP$  violation in this term. The central value obtained is  $0.0039^{+0.0036}_{-0.0032} \pm 0.0016$ . When interference is disallowed, the result becomes  $0.0021 \pm 0.0009 \pm 0.0002$ .

<sup>2</sup> This combines results of ANJOS 88c on  $K^+\pi^-$  and  $K^+\pi^-\pi^+\pi^-$  (via  $\bar{D}^0$ ) reported in the data block above (see footnotes there). It assumes no interference.

 $\Gamma(\mu^- \text{ anything via } \bar{D}^0)/\Gamma(\mu^+ \text{ anything})$  $\Gamma_{234}/\Gamma_6$ 

This is a  $D^0\text{-}\bar{D}^0$  mixing limit. See the somewhat better limits above.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.0056</b>	90	LOUIS	86 SPEC	$\pi^-$ W 225 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.012	90	BENVENUTI	85 CNTR	$\mu C$ , 200 GeV
<0.044	90	BODEK	82 SPEC	$\pi^-$ , $pFe \rightarrow D^0$

## Rare or forbidden modes

 $\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  $\Gamma_{235}/\Gamma$ 

$D^0 \rightarrow \gamma\gamma$  is a flavor-changing neutral-current decay, forbidden in the Standard Model at the tree level.

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.2</b>	90	LEES	12L BABR	$e^+e^- \approx 10.58$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 3.8	90	ABLIKIM	15F BES3	$e^+e^-$ at 3.773 GeV
<29	90	COAN	03 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

 $\Gamma(e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_{236}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.9 <math>\times 10^{-8}</math></b>	90	PETRIC	10 BELL	$e^+e^- \approx \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.7 $\times 10^{-7}$	90	LEES	12Q BABR	$e^+e^- \approx 10.58$ GeV
<1.2 $\times 10^{-6}$	90	AUBERT,B	04Y BABR	$e^+e^- \approx \mathcal{T}(4S)$
<8.19 $\times 10^{-6}$	90	PRIPSTEIN	00 E789	$p$ nucleus, 800 GeV
<6.2 $\times 10^{-6}$	90	AITALA	99G E791	$\pi^-$ N 500 GeV
<1.3 $\times 10^{-5}$	90	FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$
<1.3 $\times 10^{-4}$	90	ADLER	88 MRK3	$e^+e^-$ 3.77 GeV
<1.7 $\times 10^{-4}$	90	ALBRECHT	88G ARG	$e^+e^-$ 10 GeV
<2.2 $\times 10^{-4}$	90	HAAS	88 CLEO	$e^+e^-$ 10 GeV

 $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$  $\Gamma_{237}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.2 <math>\times 10^{-9}</math></b>	90	AAIJ	13a LHCb	$p p$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.6–8.1 $\times 10^{-7}$	90	<sup>1</sup> LEES	12Q BABR	$e^+e^- \approx 10.58$ GeV
<2.1 $\times 10^{-7}$	90	AALTONEN	10x CDF	$p\bar{p}$ , $\sqrt{s} = 1.96$ TeV
<1.4 $\times 10^{-7}$	90	PETRIC	10 BELL	$e^+e^- \approx \mathcal{T}(4S)$
<2.0 $\times 10^{-6}$	90	ABT	04 HERB	$p A$ , 920 GeV
<1.3 $\times 10^{-6}$	90	AUBERT,B	04Y BABR	$e^+e^- \approx \mathcal{T}(4S)$
<2.5 $\times 10^{-6}$	90	ACOSTA	03F CDF	See AALTONEN 10x
<1.56 $\times 10^{-5}$	90	PRIPSTEIN	00 E789	$p$ nucleus, 800 GeV
<5.2 $\times 10^{-6}$	90	AITALA	99G E791	$\pi^-$ N 500 GeV
<4.1 $\times 10^{-6}$	90	ADAMOVICH	97 BEAT	$\pi^-$ Cu, W 350 GeV
<4.2 $\times 10^{-6}$	90	ALEXOPOU...	96 E771	$p$ Si, 800 GeV
<3.4 $\times 10^{-5}$	90	FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$
<7.6 $\times 10^{-6}$	90	ADAMOVICH	95 BEAT	See ADAMOVICH 97
<4.4 $\times 10^{-5}$	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
<3.1 $\times 10^{-5}$	90	<sup>2</sup> MISHRA	94 E789	$-4.1 \pm 4.8$ events
<7.0 $\times 10^{-5}$	90	ALBRECHT	88G ARG	$e^+e^-$ 10 GeV
<1.1 $\times 10^{-5}$	90	LOUIS	86 SPEC	$\pi^-$ W 225 GeV
<3.4 $\times 10^{-4}$	90	AUBERT	85 EMC	Deep inelast. $\mu^- N$

<sup>1</sup> LEES 12Q gives a 2-sided range.

<sup>2</sup> Here MISHRA 94 uses “the statistical approach advocated by the PDG.” For an alternate approach, giving a limit of  $9 \times 10^{-6}$  at 90% confidence level, see the paper.

 $\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{238}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.5 <math>\times 10^{-5}</math></b>	90	FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

 $\Gamma(\pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{239}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.8 <math>\times 10^{-4}</math></b>	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.4 $\times 10^{-4}$	90	FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

 $\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{240}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 <math>\times 10^{-4}</math></b>	90	FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

 $\Gamma(\eta \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{241}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.3 <math>\times 10^{-4}</math></b>	90	FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

 $\Gamma(\pi^+\pi^-\pi^+e^-)/\Gamma_{\text{total}}$  $\Gamma_{242}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.73 <math>\times 10^{-4}</math></b>	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\rho^0 e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{243}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0 <math>\times 10^{-4}</math></b>	90	<sup>1</sup> FREYBERGER	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.24 $\times 10^{-4}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
<4.5 $\times 10^{-4}$	90	HAAS	88 CLEO	$e^+e^-$ 10 GeV

<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 1.8 \times 10^{-4}$  using a photon pole amplitude model.

# Meson Particle Listings

$D^0$

$\Gamma(\pi^+\pi^-\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{244}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.5 \times 10^{-7}$	90	<sup>1</sup> AAIJ	14B LHCb	$p\bar{p}$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.0 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
<sup>1</sup> AAIJ 14B measures this branching-fraction limit relative to the $\pi^+\pi^-\phi$ , $\phi \rightarrow \mu^+\mu^-$ fraction. The above limit excludes the resonant $\phi$ , $\omega$ , and $\rho$ regions, and then fills those gaps with a phase-space model.				

$\Gamma(\rho^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{245}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.2 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4.9 \times 10^{-4}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
$<2.3 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
$<8.1 \times 10^{-4}$	90	HAAS	88 CLEO	$e^+e^-$ 10 GeV
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<4.5 \times 10^{-4}$ using a photon pole amplitude model.				

$\Gamma(\omega e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{246}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-4}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<2.7 \times 10^{-4}$ using a photon pole amplitude model.				

$\Gamma(\omega\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{247}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.3 \times 10^{-4}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<6.5 \times 10^{-4}$ using a photon pole amplitude model.				

$\Gamma(K^-\bar{K}^+e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{248}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.15 \times 10^{-4}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\phi e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{249}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.2 \times 10^{-5}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5.9 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<7.6 \times 10^{-5}$ using a photon pole amplitude model.				

$\Gamma(K^-\bar{K}^+\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{250}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.3 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\phi\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{251}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4.1 \times 10^{-4}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<2.4 \times 10^{-4}$ using a photon pole amplitude model.				

$\Gamma(\bar{K}^0e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{252}/\Gamma$   
Not a useful test for  $\Delta C=1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-4}$	90	FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.7 \times 10^{-3}$	90	ADLER	89c MRK3	$e^+e^-$ 3.77 GeV

$\Gamma(\bar{K}^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{253}/\Gamma$   
Not a useful test for  $\Delta C=1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.6 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<6.7 \times 10^{-4}$	90	FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(K^-\pi^+e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{254}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.85 \times 10^{-4}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\bar{K}^*(892)^0e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{255}/\Gamma$   
Not a useful test for  $\Delta C=1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.7 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.4 \times 10^{-4}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<2.0 \times 10^{-4}$ using a photon pole amplitude model.				

$\Gamma(K^-\pi^+\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{256}/\Gamma$   
A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.59 \times 10^{-4}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\bar{K}^*(892)^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{257}/\Gamma$   
Not a useful test for  $\Delta C=1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.4 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.18 \times 10^{-3}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<1.0 \times 10^{-3}$ using a photon pole amplitude model.				

$\Gamma(\pi^+\pi^-\pi^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{258}/\Gamma$   
A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.1 \times 10^{-4}$	90	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

$\Gamma(\mu^\pm e^\mp)/\Gamma_{\text{total}}$   $\Gamma_{259}/\Gamma$   
A test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.6 \times 10^{-7}$	90	PETRIC	10 BELL	$e^+e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.3 \times 10^{-7}$	90	LEES	12Q BABR	$e^+e^- \approx 10.58$ GeV
$<8.1 \times 10^{-7}$	90	AUBERT,B	04Y BABR	$e^+e^- \approx \gamma(4S)$
$<1.72 \times 10^{-5}$	90	PRIPSTEIN	00 E789	$p$ nucleus, 800 GeV
$<8.1 \times 10^{-6}$	90	AITALA	99g E791	$\pi^- N$ 500 GeV
$<1.9 \times 10^{-5}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
$<1.0 \times 10^{-4}$	90	ALBRECHT	88g ARG	$e^+e^-$ 10 GeV
$<2.7 \times 10^{-4}$	90	HAAS	88 CLEO	$e^+e^-$ 10 GeV
$<1.2 \times 10^{-4}$	90	BECKER	87c MRK3	$e^+e^-$ 3.77 GeV
$<9 \times 10^{-4}$	90	PALKA	87 SILI	200 GeV $\pi p$
$<21 \times 10^{-4}$	90	<sup>2</sup> RILES	87 MRK2	$e^+e^-$ 29 GeV

<sup>1</sup> This is the corrected result given in the erratum to FREYBERGER 96.  
<sup>2</sup> RILES 87 assumes  $B(D \rightarrow K\pi) = 3.0\%$  and has production model dependency.

$\Gamma(\pi^0e^\pm\mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{260}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.6 \times 10^{-5}$	90	FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(\eta e^\pm\mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{261}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.0 \times 10^{-4}$	90	FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(\pi^+\pi^-\bar{e}^\pm\mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{262}/\Gamma$   
A test of lepton family-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.5 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\rho^0e^\pm\mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{263}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.9 \times 10^{-5}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<6.6 \times 10^{-5}$	90	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV
<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to $<5.0 \times 10^{-5}$ using a photon pole amplitude model.				



$\Gamma(\omega e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{264}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.2 \times 10^{-4}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The same limit is obtained using a photon pole amplitude model.

$\Gamma(K^- K^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{265}/\Gamma$   
A test of lepton family-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\phi e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{266}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-5}$	90	<sup>1</sup> FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 $<4.7 \times 10^{-5}$  90 AITALA 01c E791  $\pi^-$  nucleus, 500 GeV

<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $<3.3 \times 10^{-5}$  using a photon pole amplitude model.

$\Gamma(K^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{267}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.0 \times 10^{-4}$	90	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

$\Gamma(K^- \pi^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{268}/\Gamma$   
A test of lepton family-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.53 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\bar{K}^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{269}/\Gamma$   
A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.3 \times 10^{-5}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.0 \times 10^{-4}$  90 <sup>1</sup> FREYBERGER 96 CLE2  $e^+ e^- \approx \gamma(4S)$

<sup>1</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The same limit is obtained using a photon pole amplitude model.

$\Gamma(2\pi^- 2e^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{270}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.12 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(2\pi^- 2\mu^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{271}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.9 \times 10^{-5}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(K^- \pi^- 2e^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{272}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.06 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(K^- \pi^- 2\mu^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{273}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.9 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(2K^- 2e^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{274}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.52 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(2K^- 2\mu^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{275}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<9.4 \times 10^{-5}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\pi^- \pi^- e^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{276}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.9 \times 10^{-5}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(K^- \pi^- e^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{277}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.18 \times 10^{-4}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(2K^- e^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{278}/\Gamma$   
A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.7 \times 10^{-5}$	90	AITALA 01c	E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(p e^-)/\Gamma_{\text{total}}$   $\Gamma_{279}/\Gamma$   
A test of baryon- and lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.0 \times 10^{-5}$	90	<sup>1</sup> RUBIN 09	CLEO	$e^+ e^-$ at $\psi(3770)$

<sup>1</sup> This RUBIN 09 limit is for either  $D^0 \rightarrow p e^-$  or  $\bar{D}^0 \rightarrow p e^-$  decay.

$\Gamma(\bar{p} e^-)/\Gamma_{\text{total}}$   $\Gamma_{280}/\Gamma$   
A test of baryon- and lepton-number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-5}$	90	<sup>1</sup> RUBIN 09	CLEO	$e^+ e^-$ at $\psi(3770)$

<sup>1</sup> This RUBIN 09 limit is for either  $D^0 \rightarrow \bar{p} e^-$  or  $\bar{D}^0 \rightarrow \bar{p} e^-$  decay.

### $D^0$ CP-VIOLATING DECAY-RATE ASYMMETRIES

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for the decay to state  $f$ , divided by the sum of the widths:

$$A_{CP}(f) = [\Gamma(D^0 \rightarrow f) - \Gamma(\bar{D}^0 \rightarrow \bar{f})] / [\Gamma(D^0 \rightarrow f) + \Gamma(\bar{D}^0 \rightarrow \bar{f})].$$

$A_{CP}(K^+ K^-)$  in  $D^0, \bar{D}^0 \rightarrow K^+ K^-$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.14 \pm 0.12</math> OUR AVERAGE</b>				
$-0.06 \pm 0.15 \pm 0.10$	1.8M	<sup>1</sup> AAIJ	14AK	LHCB Time-integrated
$-0.24 \pm 0.22 \pm 0.09$	476k	<sup>1</sup> AALTONEN	12B	CDF $p\bar{p}$ , $\sqrt{s}=1.96$ TeV
$0.00 \pm 0.34 \pm 0.13$	129k	<sup>2</sup> AUBERT	08M	BABR $e^+ e^- \approx 10.6$ GeV
$-0.43 \pm 0.30 \pm 0.11$	120k	<sup>3</sup> STARIC	08	BELL $e^+ e^- \approx \gamma(4S)$
$+2.0 \pm 1.2 \pm 0.6$		<sup>4</sup> ACOSTA	05c	CDF $p\bar{p}$ , $\sqrt{s}=1.96$ TeV
$0.0 \pm 2.2 \pm 0.8$	3023	<sup>4</sup> CSORNA	02	CLE2 $e^+ e^- \approx \gamma(4S)$
$-0.1 \pm 2.2 \pm 1.5$	3330	<sup>4</sup> LINK	00B	FOCS
$-1.0 \pm 4.9 \pm 1.2$	609	<sup>4</sup> AITALA	98c	E791 $-0.093 < A_{CP} < +0.073$ (90% CL)

<sup>1</sup> See also " $D^0$  CP-violating asymmetry differences" at the end of the CP-violating asymmetries.

<sup>2</sup> AUBERT 08M uses corrected numbers of events directly, not ratios with  $K^\mp \pi^\pm$  events.

<sup>3</sup> STARIC 08 uses  $D^0 \rightarrow K^- \pi^+$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  decays to correct for detector-induced asymmetries.

<sup>4</sup> AITALA 98c, LINK 00B, CSORNA 02, and ACOSTA 05c measure  $N(D^0 \rightarrow K^+ K^-)/N(D^0 \rightarrow K^- \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $\bar{D}^0$ .

$A_{CP}(K_S^0 K_S^0)$  in  $D^0, \bar{D}^0 \rightarrow K_S^0 K_S^0$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-5 \pm 5</math> OUR AVERAGE</b>				
$-2.9 \pm 5.2 \pm 2.2$	630	AAIJ	15AT	LHCB $p\bar{p}$ at 7, 8 TeV
$-23 \pm 19$	65	BONVICINI	01	CLE2 $e^+ e^- \approx 10.6$ GeV

$A_{CP}(\pi^+ \pi^-)$  in  $D^0, \bar{D}^0 \rightarrow \pi^+ \pi^-$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.15</math> OUR AVERAGE</b>				
$-0.20 \pm 0.19 \pm 0.10$	774k	<sup>1,2</sup> AAIJ	14AK	LHCB Time-integrated
$0.22 \pm 0.24 \pm 0.11$	215k	<sup>1</sup> AALTONEN	12B	CDF $p\bar{p}$ , $\sqrt{s}=1.96$ TeV
$-0.24 \pm 0.52 \pm 0.22$	63.7k	<sup>3</sup> AUBERT	08M	BABR $e^+ e^- \approx 10.6$ GeV
$0.43 \pm 0.52 \pm 0.12$	51k	<sup>4</sup> STARIC	08	BELL $e^+ e^- \approx \gamma(4S)$
$1.0 \pm 1.3 \pm 0.6$		<sup>5</sup> ACOSTA	05c	CDF $p\bar{p}$ , $\sqrt{s}=1.96$ TeV
$1.9 \pm 3.2 \pm 0.8$	1136	<sup>5</sup> CSORNA	02	CLE2 $e^+ e^- \approx \gamma(4S)$
$4.8 \pm 3.9 \pm 2.5$	1177	<sup>5</sup> LINK	00B	FOCS
$-4.9 \pm 7.8 \pm 3.0$	343	<sup>5</sup> AITALA	98c	E791 $-0.186 < A_{CP} < +0.088$ (90% CL)

<sup>1</sup> See also " $D^0$  CP-violating asymmetry differences" at the end of the CP-violating asymmetries.

<sup>2</sup> AAIJ 14AK uses  $\Delta A_{CP}(\pi\pi, KK)$  and  $A_{CP}(KK)$  reported in the same paper.

<sup>3</sup> AUBERT 08M uses corrected numbers of events directly, not ratios with  $K^\mp \pi^\pm$  events.

<sup>4</sup> STARIC 08 uses  $D^0 \rightarrow K^- \pi^+$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  decays to correct for detector-induced asymmetries.

<sup>5</sup> AITALA 98c, LINK 00B, CSORNA 02, and ACOSTA 05c measure  $N(D^0 \rightarrow \pi^+ \pi^-)/N(D^0 \rightarrow K^- \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $\bar{D}^0$ .

$A_{CP}(\pi^0 \pi^0)$  in  $D^0, \bar{D}^0 \rightarrow \pi^0 \pi^0$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.0 \pm 0.6</math> OUR AVERAGE</b>				
$-0.03 \pm 0.64 \pm 0.10$	34k	NISAR	14	BELL $e^+ e^-$ at/near $\gamma$ 's
$0.1 \pm 4.8$	810	BONVICINI	01	CLE2 $e^+ e^- \approx 10.6$ GeV

# Meson Particle Listings

$D^0$

$A_{CP}(\pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \pi^+\pi^-\pi^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3 ± 0.4 OUR AVERAGE</b>				
0.43 ± 1.30	123k ± 490	ARINSTEIN 08	BELL	$e^+e^- \approx \Upsilon(4S)$
0.31 ± 0.41 ± 0.17	80 ± .3k	<sup>1</sup> AUBERT 08A0	BABR	$e^+e^- \approx 10.6$ GeV
1 $\begin{smallmatrix} +9 \\ -7 \end{smallmatrix} \pm 5$		CRONIN-HEN..05	CLEO	$e^+e^- \approx 10$ GeV

<sup>1</sup> AUBERT 08A0 report their result using a different sign convention.

$A_{CP}(\rho(770)^+\pi^- \rightarrow \pi^+\pi^-\pi^0)$ in $D^0 \rightarrow \rho^+\pi^-, \bar{D}^0 \rightarrow \rho^-\pi^+$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+1.2 ± 0.8 ± 0.3</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(770)^0\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \rho^0\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−3.1 ± 2.7 ± 1.2</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(770)^-\pi^+ \rightarrow \pi^+\pi^-\pi^0)$ in $D^0 \rightarrow \rho^-\pi^+, \bar{D}^0 \rightarrow \rho^+\pi^-$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−1.0 ± 1.6 ± 0.7</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(1450)^+\pi^- \rightarrow \pi^+\pi^-\pi^0)$ in $D^0 \rightarrow \rho(1450)^+\pi^-, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>0 ± 50 ± 50</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(1450)^0\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \rho(1450)^0\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−17 ± 33 ± 17</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(1450)^-\pi^+ \rightarrow \pi^+\pi^-\pi^0)$ in $D^0 \rightarrow \rho(1450)^-\pi^+, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+6 ± 8 ± 3</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(1700)^+\pi^- \rightarrow \pi^+\pi^-\pi^0)$ in $D^0 \rightarrow \rho(1700)^+\pi^-, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−5 ± 13 ± 5</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(1700)^0\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \rho(1700)^0\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+13 ± 8 ± 3</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\rho(1700)^-\pi^+ \rightarrow \pi^+\pi^-\pi^0)$ in $D^0 \rightarrow \rho(1700)^-\pi^+, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+8 ± 10 ± 5</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(f_0(980)\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_0(980)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>0 ± 25 ± 25</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(f_0(1370)\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_0(1370)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+25 ± 13 ± 13</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(f_0(1500)\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_0(1500)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>0 ± 13 ± 13</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(f_0(1710)\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_0(1710)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>0 ± 17 ± 17</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(f_2(1270)\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_2(1270)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−4 ± 4 ± 4</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\sigma(400)\pi^0 \rightarrow \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \sigma(400)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+6 ± 6 ± 6</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\text{nonresonant } \pi^+\pi^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \text{nonresonant } \pi^+\pi^-\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−13 ± 19 ± 13</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(2\pi^+2\pi^-)$ in $D^0, \bar{D}^0 \rightarrow 2\pi^+2\pi^-$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>no evidence</b>	<sup>1</sup> AAIJ	13BR LHCB		

<sup>1</sup> AAIJ 13BR searched for  $CP$  violation in binned phase space. No evidence was found.

$A_{CP}(K^+K^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow K^+K^-\pi^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>−1.00 ± 1.67 ± 0.25</b>	11 ± 0.11k	AUBERT	08A0 BABR	$e^+e^- \approx 10.6$ GeV

$A_{CP}(K^*(892)^+K^- \rightarrow K^+K^-\pi^0)$ in $D^0 \rightarrow K^*(892)^+K^-, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−0.9 ± 1.2 ± 0.4</b>	<sup>1</sup> AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

<sup>1</sup> AUBERT 08A0 report their result using a different sign convention.

$A_{CP}(K^*(1410)^+K^- \rightarrow K^+K^-\pi^0)$ in $D^0 \rightarrow K^*(1410)^+K^-, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−21 ± 23 ± 8</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}((K^+\pi^0)_S\text{--}wave K^- \rightarrow K^+K^-\pi^0)$ in $D^0 \rightarrow (K^+\pi^0)_S K^-, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+7 ± 15 ± 3</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(\phi(1020)\pi^0 \rightarrow K^+K^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow \phi(1020)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>+1.1 ± 2.1 ± 0.5</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(f_0(980)\pi^0 \rightarrow K^+K^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_0(980)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−3 ± 19 ± 1</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(a_0(980)^0\pi^0 \rightarrow K^+K^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow a_0(980)^0\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−5 ± 16 ± 2</b>	<sup>1</sup> AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

<sup>1</sup> This AUBERT 08A0 value is obtained when the  $a_0(980)^0$  replaces the  $f_0(980)$  in the fit.

$A_{CP}(f_2'(1525)\pi^0 \rightarrow K^+K^-\pi^0)$ in $D^0, \bar{D}^0 \rightarrow f_2'(1525)\pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>0 ± 50 ± 150</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(K^*(892)^-K^+ \rightarrow K^+K^-\pi^0)$ in $D^0 \rightarrow K^*(892)^-K^+, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−5 ± 4 ± 1</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(K^*(1410)^-K^+ \rightarrow K^+K^-\pi^0)$ in $D^0 \rightarrow K^*(1410)^-K^+, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−17 ± 28 ± 7</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}((K^-\pi^0)_S\text{--}wave K^+ \rightarrow K^+K^-\pi^0)$ in $D^0 \rightarrow (K^-\pi^0)_S K^+, \bar{D}^0 \rightarrow$ c.c.				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−7 ± 40 ± 8</b>	AUBERT	08A0 BABR	Table 1, −Col.5/2×Col.2	

$A_{CP}(K_S^0\pi^0)$ in $D^0, \bar{D}^0 \rightarrow K_S^0\pi^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>−0.20 ± 0.17 OUR AVERAGE</b>				
−0.21 ± 0.16 ± 0.07	467k	<sup>1</sup> NISAR	14	BELL $e^+e^-$ at/near $\Upsilon$ 's
0.1 ± 1.3	9099	BONVICINI	01	CLE2 $e^+e^- \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.28 ± 0.19 ± 0.10	326k	KO	11	BELL See NISAR 14
−1.8 ± 3.0		BARTELT	95	CLE2 See BONVICINI 01

<sup>1</sup> After subtracting CPV in  $K^0 - \bar{K}^0$  mixing, NISAR 14 gets  $A_{CP} = (+0.12 \pm 0.16 \pm 0.07)\%$ .

$A_{CP}(K_S^0\eta)$ in $D^0, \bar{D}^0 \rightarrow K_S^0\eta$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.54 ± 0.51 ± 0.16</b>	46k	KO	11	BELL $e^+e^- \approx \Upsilon(4S)$

$A_{CP}(K_S^0\eta')$ in $D^0, \bar{D}^0 \rightarrow K_S^0\eta'$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.98 ± 0.67 ± 0.14</b>	27k	KO	11	BELL $e^+e^- \approx \Upsilon(4S)$

$A_{CP}(K_S^0\phi)$ in $D^0, \bar{D}^0 \rightarrow K_S^0\phi$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>−2.8 ± 9.4</b>	BARTELT	95	CLE2	−18.2 < $A_{CP}$ < +12.6% (90%CL)

$A_{CP}(K^\mp\pi^\pm)$ in $D^0 \rightarrow K^-\pi^+, \bar{D}^0 \rightarrow K^+\pi^-$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3 ± 0.3 ± 0.6</b>		BONVICINI	14	CLEO All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
+0.5 ± 0.4 ± 0.9	150k	MENDEZ	10	CLEO See BONVICINI 14
−0.4 ± 0.5 ± 0.9		DOBBS	07	CLEO See BONVICINI 14

$A_{CP}(K^\pm\pi^\mp)$ in $D^0 \rightarrow K^+\pi^-, \bar{D}^0 \rightarrow K^-\pi^+$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0 ± 1.6 OUR AVERAGE</b>				
−0.7 ± 1.9		<sup>1</sup> AAIJ	13CE	LHCB $p\bar{p}$ at 7, 8 TeV
−2.1 ± 5.2 ± 1.5	4.0k	AUBERT	07W	BABR $e^+e^- \approx 10.6$ GeV
+2.3 ± 4.7	4.0k	<sup>2</sup> ZHANG	06	BELL $e^+e^-$
+18 ± 14 ± 4		<sup>3</sup> LINK	05H	FOCS $\gamma$ nucleus
+9.5 ± 6.1 ± 8.3		<sup>4</sup> AUBERT	03Z	BABR $e^+e^-$ , 10.6 GeV
+2 $\begin{smallmatrix} +19 \\ -20 \end{smallmatrix} \pm 1$	45	<sup>5</sup> GODANG	00	CLE2 $e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−8.0 ± 7.7	0.8k	<sup>6</sup> LI	05A	BELL See ZHANG 06

<sup>1</sup> Based on 3 fb<sup>-1</sup> of data collected at  $\sqrt{s} = 7, 8$  TeV. Allowing for  $CP$  violation, the direct  $CP$ -violation in mixing is reported for the  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$ .

<sup>2</sup> This ZHANG 06 result allows mixing.

<sup>3</sup> This LINK 05H result assumes no mixing. If mixing is allowed, it becomes  $0.13^{+0.33}_{-0.25} \pm 0.10$ .

<sup>4</sup> This AUBERT 03Z limit assumes no mixing. If mixing is allowed, the 95% confidence-level interval is  $(-2.8 < A_D < 4.9) \times 10^{-3}$ .

<sup>5</sup> This GODANG 00 result assumes no  $D^0$ - $\bar{D}^0$  mixing and becomes  $-0.43 < A_{CP} < +0.34$  at 95% CL. If mixing is allowed  $A_{CP} = -0.01^{+0.16}_{-0.17} \pm 0.01$ .

<sup>6</sup> This LI 05A result allows mixing.

### $A_{CP}(K^- \pi^+) \text{ in } D_{CP}(\pm 1) \rightarrow K^\mp \pi^\pm$

$$A_{CP}(K^- \pi^+) = [B(D_{CP}(-) \rightarrow K^- \pi^+ + \text{c.c.}) - B(D_{CP}(+) \rightarrow K^- \pi^+ + \text{c.c.})] / \text{Sum.}$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>12.7 ± 1.3 ± 0.7</b>	<sup>1</sup> ABLIKIM	14c BES3	$e^+ e^- \rightarrow D^0 \bar{D}^0$ , 3.77 GeV
<sup>1</sup> ABLIKIM 14c uses quantum correlations in $e^+ e^- \rightarrow D^0 \bar{D}^0$ at the $\psi(3770)$ to measure the asymmetry of the branching fraction of $D^0 \rightarrow K^- \pi^+$ in $CP$ -odd and $CP$ -even eigenstates. It then extracts the strong-phase difference $\delta_{K\pi}$ .			

### $A_{CP}(K^\mp \pi^\pm \pi^0) \text{ in } D^0 \rightarrow K^- \pi^+ \pi^0, \bar{D}^0 \rightarrow K^+ \pi^- \pi^0$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.1 ± 0.5 OUR AVERAGE</b>			
0.1 ± 0.3 ± 0.4	BONVICINI	14 CLEO	All CLEO-c runs
-1 ± 8.6	<sup>1</sup> KOPP	01 CLE2	$e^+ e^- \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.2 ± 0.4 ± 0.8	DOBBS	07 CLEO	See BONVICINI 14
<sup>1</sup> KOPP 01 fits separately the $D^0$ and $\bar{D}^0$ Dalitz plots and then calculates the integrated difference of normalized densities divided by the integrated sum.			

### $A_{CP}(K^\pm \pi^\mp \pi^0) \text{ in } D^0 \rightarrow K^+ \pi^- \pi^0, \bar{D}^0 \rightarrow K^- \pi^+ \pi^0$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0 ± 5 OUR AVERAGE</b>				
-0.6 ± 5.3	1978 ± 104	TIAN	05 BELL	$e^+ e^- \approx 7(45)$
+9 ± 25	38	BRANDENB...	01 CLE2	$e^+ e^- \approx 7(45)$

### $A_{CP}(K_S^0 \pi^+ \pi^-) \text{ in } D^0, \bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.1 ± 0.8 OUR AVERAGE</b>				
-0.05 ± 0.57 ± 0.54	350k	<sup>1</sup> AALTONEN	12AD CDF	
-0.9 ± 2.1 ± 1.6	4854	<sup>2</sup> ASNER	04A CLEO	$e^+ e^- \approx 10$ GeV

<sup>1</sup> This is the overall result of AALTONEN 12AD. Following are the 15  $CP$  fit-fraction asymmetries from the amplitude analysis of the  $D^0$  and  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  Dalitz plots.

<sup>2</sup> This is the overall result of ASNER 04A;  $CP$ -violating limits are also given below for each of the 10 resonant submodes found in an amplitude analysis of the  $D^0$  and  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  Dalitz plots.

### $A_{CP}(K^*(892)^\mp \pi^\pm \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow K^* \pi^+, \bar{D}^0 \rightarrow K^* \pi^-$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>+0.36 ± 0.33 ± 0.40</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+2.5 ± 1.9 ± 3.3	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K^*(892)^\pm \pi^\mp \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow K^* \pi^-, \bar{D}^0 \rightarrow K^* \pi^+$

This is a doubly Cabibbo-suppressed mode.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>+1.0 ± 5.7 ± 2.1</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-21 ± 42 ± 28	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_S^0 \rho^0 \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow \bar{K}^0 \rho^0, \bar{D}^0 \rightarrow K^0 \rho^0$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-0.05 ± 0.50 ± 0.08</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+3.1 ± 3.8 ± 2.7	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_S^0 \omega \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow \bar{K}^0 \omega, \bar{D}^0 \rightarrow K^0 \omega$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-12.6 ± 6.0 ± 2.6</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-26 ± 24 ± 22	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_S^0 f_0(980) \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow \bar{K}^0 f_0(980), \bar{D}^0 \rightarrow K^0 f_0(980)$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-0.4 ± 2.2 ± 1.6</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-4.7 ± 11.0 ± 24.9	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_S^0 f_2(1270) \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow \bar{K}^0 f_2(1270), \bar{D}^0 \rightarrow K^0 f_2(1270)$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-4.0 ± 3.4 ± 3.0</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+34 ± 51 ± 33	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_S^0 f_0(1370) \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow \bar{K}^0 f_0(1370), \bar{D}^0 \rightarrow K^0 f_0(1370)$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-0.5 ± 4.6 ± 7.7</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+18 ± 10 ± 13	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_S^0 \rho^0(1450) \text{ in } D^0 \rightarrow \bar{K}^0 \rho^0(1450), \bar{D}^0 \rightarrow K^0 \rho^0(1450)$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-4.1 ± 5.2 ± 8.1</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts

### $A_{CP}(K_S^0 f_0(600) \text{ in } D^0 \rightarrow \bar{K}^0 f_0(600), \bar{D}^0 \rightarrow K^0 f_0(600)$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-2.7 ± 2.7 ± 3.6</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts

### $A_{CP}(K^*(1410)^\mp \pi^\pm) \text{ in } D^0 \rightarrow K^*(1410)^- \pi^+, \bar{D}^0 \rightarrow K^*(1410)^+ \pi^-$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-2.3 ± 5.7 ± 6.4</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts

### $A_{CP}(K_0^*(1430)^\mp \pi^\pm \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow K_0^*(1430)^- \pi^+, \bar{D}^0 \rightarrow \text{c.c.}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>4.0 ± 2.4 ± 3.8</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.2 ± 11.3 ± 8.8	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_0^*(1430)^\pm \pi^\mp) \text{ in } D^0 \rightarrow K_0^*(1430)^+ \pi^-, \bar{D}^0 \rightarrow K_0^*(1430)^- \pi^+$

This is a doubly Cabibbo-suppressed mode.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>+12 ± 11 ± 10</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts

### $A_{CP}(K_2^*(1430)^\mp \pi^\pm \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow K_2^*(1430)^- \pi^+, \bar{D}^0 \rightarrow \text{c.c.}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>+2.9 ± 4.0 ± 4.1</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-7 ± 25 ± 13	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K_2^*(1430)^\pm \pi^\mp) \text{ in } D^0 \rightarrow K_2^*(1430)^+ \pi^-, \bar{D}^0 \rightarrow K_2^*(1430)^- \pi^+$

This is a doubly Cabibbo-suppressed mode.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-10 ± 14 ± 29</b>	AALTONEN	12AD CDF	Dalitz fit, ~ 350k evts

### $A_{CP}(K^*(1680)^\mp \pi^\pm \rightarrow K_S^0 \pi^+ \pi^-) \text{ in } D^0 \rightarrow K^*(1680)^- \pi^+, \bar{D}^0 \rightarrow \text{c.c.}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-36 ± 19 ± 10	ASNER	04A CLEO	Dalitz fit, 4854 evts

### $A_{CP}(K^- \pi^+ \pi^+ \pi^-) \text{ in } D^0 \rightarrow K^- \pi^+ \pi^+ \pi^-, \bar{D}^0 \rightarrow K^+ \pi^- \pi^- \pi^+$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.2 ± 0.3 ± 0.4</b>	BONVICINI	14 CLEO	All CLEO-c runs
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+0.7 ± 0.5 ± 0.9	DOBBS	07 CLEO	See BONVICINI 14

### $A_{CP}(K^\pm \pi^\mp \pi^+ \pi^-) \text{ in } D^0 \rightarrow K^+ \pi^- \pi^+ \pi^-, \bar{D}^0 \rightarrow K^- \pi^+ \pi^+ \pi^-$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-1.8 ± 4.4</b>	1721 ± 75	TIAN	05 BELL	$e^+ e^- \approx 7(45)$

### $A_{CP}(K^+ K^- \pi^+ \pi^-) \text{ in } D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^+ \pi^-$

See also AAIJ 13BR for a search for  $CP$  violation in  $D^0 \rightarrow K^+ K^- \pi^+ \pi^-$  in binned phase space. No evidence of  $CP$  violation was found.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-8.2 ± 5.6 ± 4.7</b>	828 ± 46	LINK	05E FOCS	$\gamma A, \bar{E}_{\gamma} \approx 180$ GeV

### $A_{CP}(K_1^*(1270)^+ K^- \rightarrow K^{*0} \pi^+ K^-) \text{ in } D^0 \rightarrow K_1^*(1270)^+ K^-, \bar{D}^0 \rightarrow \text{c.c.}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-0.7 ± 10.4</b>	ARTUSO	12 CLEO	Amplitude fit, 2959 evts.

### $A_{CP}(K_1^*(1270)^- K^+ \rightarrow \bar{K}^{*0} \pi^- K^+) \text{ in } D^0 \rightarrow K_1^*(1270)^- K^+, \bar{D}^0 \rightarrow \text{c.c.}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-10.0 ± 31.5</b>	ARTUSO	12 CLEO	Amplitude fit, 2959 evts.

### $A_{CP}(K_1^*(1270)^+ K^- \rightarrow \rho^0 K^+ K^-) \text{ in } D^0 \rightarrow K_1^*(1270)^+ K^-, \bar{D}^0 \rightarrow \text{c.c.}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-6.5 ± 16.9</b>	ARTUSO	12 CLEO	Amplitude fit, 2959 evts.

# Meson Particle Listings

## $D^0$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$+9.6 \pm 12.9$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-20.0 \pm 16.8$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-1.1 \pm 13.7$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$+9.5 \pm 13.5$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-2.7 \pm 5.3$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-37.1 \pm 19.0$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$-8.6 \pm 10.4$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$+2.7 \pm 10.6$	ARTUSO	12	CLEO Amplitude fit, 2959 evts.

### $D^0$ CP-EVEN FRACTIONS

The  $CP$ -even fraction  $F_+$ , defined for self-conjugate final states, like the coherence factor is useful for measuring the unitary triangle angle  $\gamma$  in  $B \rightarrow DK$  decays. A purely  $CP$ -even state has  $F_+ = 1$ , a  $CP$ -odd one has  $F_+ = 0$ . For details, see NAYAK 15.

VALUE (%)	DOCUMENT ID	COMMENT
$97.3 \pm 1.7$	MALDE	15 Uses CLEO data
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$96.8 \pm 1.7 \pm 0.6$	NAYAK	15 see MALDE 15

VALUE (%)	DOCUMENT ID	COMMENT
$73.2 \pm 5.5$	MALDE	15 Uses CLEO data
• • • We do not use the following data for averages, fits, limits, etc. • • •		
$73.1 \pm 5.8 \pm 2.1$	NAYAK	15 see MALDE 15

VALUE (%)	DOCUMENT ID	COMMENT
$73.7 \pm 2.8$	MALDE	15 Uses CLEO data

### $D^0$ CP-VIOLATING ASYMMETRY DIFFERENCES

$$\Delta A_{CP} = A_{CP}(K^+ K^-) - A_{CP}(\pi^+ \pi^-)$$

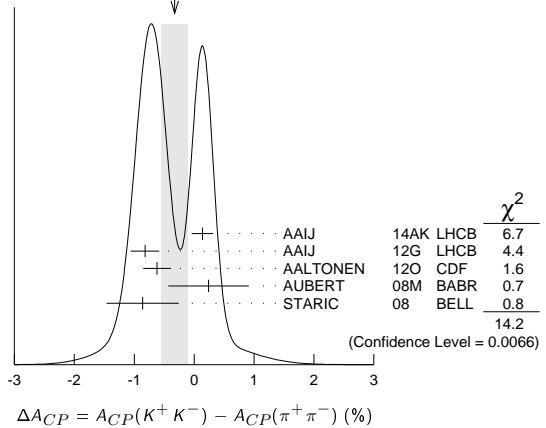
$CP$  violation in these modes can come from the decay amplitudes (direct) and/or from mixing or interference of mixing and decay (indirect). The difference  $\Delta A_{CP}$  is primarily sensitive to the direct component, and only retains a second-order dependence on the indirect component for measurements where the mean decay time of the  $K^+ K^-$  and  $\pi^+ \pi^-$  samples are not identical. The results below are averaged assuming the indirect component can be neglected.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
$-0.32 \pm 0.22$ OUR AVERAGE	Error includes scale factor of 1.9. See the ideogram below.			
$-0.14 \pm 0.16 \pm 0.08$	2.17/0.77M	<sup>1</sup> AAIJ	14AK LHCB	Time-integrated
$-0.82 \pm 0.21 \pm 0.11$		<sup>1</sup> AAIJ	12G LHCB	Time-integrated
$-0.62 \pm 0.21 \pm 0.10$		<sup>1</sup> AALTONEN	120 CDF	Time-integrated
$0.24 \pm 0.62 \pm 0.26$		<sup>1,2</sup> AUBERT	08M BABR	Time-integrated
$-0.86 \pm 0.60 \pm 0.07$	120k	<sup>1</sup> STARIC	08 BELL	Time-integrated
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.49 \pm 0.30 \pm 0.14$	559/222k	AAIJ	13AD LHCB	See AAIJ 14AK
$-0.46 \pm 0.31 \pm 0.12$		AALTONEN	12B CDF	See AALTONEN 120

<sup>1</sup>  $D$ 's from  $D^*$ 's decays.

<sup>2</sup> Calculated from the AUBERT 08M values of  $A_{CP}(K^+ K^-)$  and  $A_{CP}(\pi^+ \pi^-)$ . The systematic error here combines the systematic errors in quadrature, and therefore somewhat over-estimates it.

WEIGHTED AVERAGE  
-0.32±0.22 (Error scaled by 1.9)



### $D^0$ $\chi^2$ TESTS OF $CP$ -VIOLATION ( $CPV$ )

We list model-independent searches for local  $CP$  violation in phase-space distributions of multi-body decays.

Most of these searches divide phase space (Dalitz plot for 3-body decays, five-dimensional equivalent for 4-body decays) into bins, and perform a  $\chi^2$  test comparing normalised yields  $N_i, \bar{N}_i$  in  $CP$ -conjugate bin pairs  $i$ :  $\chi^2 = \sum_i (N_i - \alpha \bar{N}_i) / \sigma(N_i - \alpha \bar{N}_i)$ . The factor  $\alpha = (\sum_i N_i) / (\sum_i \bar{N}_i)$  removes the dependence on phase-space-integrated rate asymmetries. The result is used to obtain the probability (p-value) to obtain the measured  $\chi^2$  or larger under the assumption of  $CP$  conservation [AUBERT 08A0, BEDIAGA 09]. Alternative methods obtain p-values from other test variables based on unbinned analyses [WILLIAMS 11, AAJ 14C]. Results can be combined using Fisher's method [MOSTELLER 48].

Local $CPV$ in $D^0, \bar{D}^0 \rightarrow \pi^+ \pi^- \pi^0$	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.9 OUR EVALUATION</b>				
2.6	566k	<sup>1</sup> AAIJ	15A LHCB	unbinned method
32.8	82k	AUBERT	08A0 BABR	$\chi^2$
<sup>1</sup> Unusually, AAJ 15A assigns an uncertainty on the p value of $\pm 0.5\%$ . This results from limited test statistics.				

Local $CPV$ in $D^0, \bar{D}^0 \rightarrow \pi^+ \pi^- \pi^+ \pi^-$	EVTS	DOCUMENT ID	TECN	COMMENT
<b>41</b>	330k	AAIJ	13BR LHCB	$\chi^2$

Local $CPV$ in $D^0, \bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$	EVTS	DOCUMENT ID	TECN	COMMENT
<b>96</b>	350k	AALTONEN	12AD CDF	$\chi^2$

Local $CPV$ in $D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^0$	EVTS	DOCUMENT ID	TECN	COMMENT
<b>16.6</b>	11k	AUBERT	08A0 BABR	$\chi^2$

Local $CPV$ in $D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^+ \pi^-$	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.1</b>	57k	AAIJ	13BR LHCB	$\chi^2$

### $CP$ VIOLATING ASYMMETRIES OF $P$ -ODD ( $T$ -ODD) MOMENTS

The  $CP$ -sensitive  $P$ -odd ( $T$ -odd) correlation in  $D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^+ \pi^-$  decays.  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0 \pi^+$  and  $D^{*-} \rightarrow \bar{D}^0 \pi^-$ .

$A_{Tviol}(K^+ K^- \pi^+ \pi^-)$ in $D^0, \bar{D}^0 \rightarrow K^+ K^- \pi^+ \pi^-$	EVTS	DOCUMENT ID	TECN	COMMENT
$C_T \equiv \bar{p}_{K^+} \cdot (\bar{p}_{\pi^+} \times \bar{p}_{\pi^-})$ is a parity-odd correlation of the $K^+$ , $\pi^+$ , and $\pi^-$ momenta (evaluated in the $D^0$ rest frame) for the $D^0$ . $\bar{C}_T \equiv \bar{p}_{K^-} \cdot (\bar{p}_{\pi^-} \times \bar{p}_{\pi^+})$ is the corresponding quantity for the $\bar{D}^0$ . Then				
$A_T \equiv [\Gamma(C_T > 0) - \Gamma(C_T < 0)] / [\Gamma(C_T > 0) + \Gamma(C_T < 0)]$ , and				
$\bar{A}_T \equiv [\Gamma(-\bar{C}_T > 0) - \Gamma(-\bar{C}_T < 0)] / [\Gamma(-\bar{C}_T > 0) + \Gamma(-\bar{C}_T < 0)]$ , and				
$A_{Tviol} \equiv \frac{1}{2}(A_T - \bar{A}_T)$ . $C_T$ and $\bar{C}_T$ are commonly referred to as $T$ -odd moments, because they are odd under $T$ reversal. However, the $T$ -conjugate process $K^+ K^- \pi^+ \pi^- \rightarrow D^0$ is not accessible, while the $P$ -conjugate process is.				

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.7 ± 2.7 OUR AVERAGE</b>				
$1.8 \pm 2.9 \pm 0.4$	171k	AAIJ	14BC LHCB	$B \rightarrow D^0 \mu^- X$
$1.0 \pm 5.1 \pm 4.4$	47k	DEL-AMO-SA.10	BABR	$e^+ e^- \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$10 \pm 57 \pm 37$	0.8k	LINK	05E FOCS	$\gamma A, \bar{E}_\gamma \approx 180$ GeV

$D^0$  CPT-VIOLATING DECAY-RATE ASYMMETRIES $A_{CP\mathcal{T}}(K^\mp\pi^\pm)$  in  $D^0 \rightarrow K^\mp\pi^\pm$ ,  $\bar{D}^0 \rightarrow K^\pm\pi^\mp$ 

$A_{CP\mathcal{T}}(t)$  is defined in terms of the time-dependent decay probabilities  $P(D^0 \rightarrow K^\mp\pi^\pm)$  and  $\bar{P}(\bar{D}^0 \rightarrow K^\pm\pi^\mp)$  by  $A_{CP\mathcal{T}}(t) = (\bar{P} - P)/(\bar{P} + P)$ . For small mixing parameters  $x \equiv \Delta m/\Gamma$  and  $y \equiv \Delta\Gamma/2\Gamma$  (as is the case), and times  $t$ ,  $A_{CP\mathcal{T}}(t)$  reduces to  $[y \operatorname{Re} \xi - x \operatorname{Im} \xi] \Gamma t$ , where  $\xi$  is the CPT-violating parameter.

The following is actually  $y \operatorname{Re} \xi - x \operatorname{Im} \xi$ .

VALUE	DOCUMENT ID	TECN	COMMENT
$0.0083 \pm 0.0065 \pm 0.0041$	LINK	03B	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

 $D^0 \rightarrow K^*(892)^-\ell^+\nu_\ell$  FORM FACTORS $r_V \equiv V(0)/A_1(0)$  in  $D^0 \rightarrow K^*(892)^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$1.71 \pm 0.68 \pm 0.34$	LINK	05B	FOCS $K^*(892)^-\mu^+\nu_\mu$

 $r_2 \equiv A_2(0)/A_1(0)$  in  $D^0 \rightarrow K^*(892)^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.91 \pm 0.37 \pm 0.10$	LINK	05B	FOCS $K^*(892)^-\mu^+\nu_\mu$

 $D^0 \rightarrow K^-/\pi^-\ell^+\nu_\ell$  FORM FACTORS $f_+(0)$  in  $D^0 \rightarrow K^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.736 \pm 0.004$ OUR AVERAGE			
$0.7368 \pm 0.0026 \pm 0.0036$	71k	ABLIKIM	15X BES3 $\ell=e$ , 2-parameter fit
$0.727 \pm 0.007 \pm 0.009$		AUBERT	07Bg BABR $\ell=e$ , 2-parameter fit

 $f_+(0)|V_{cs}|$  in  $D^0 \rightarrow K^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.719 \pm 0.004$ OUR AVERAGE			
$0.7172 \pm 0.0025 \pm 0.0035$	71k	1 ABLIKIM	15X BES3 $\ell=e$ , 2-parameter fit
$0.726 \pm 0.008 \pm 0.004$		BESSION	09 CLEO $\ell=e$ , 3-parameter fit

<sup>1</sup> The 3-parameter fit yields  $0.7195 \pm 0.0035 \pm 0.0041$ .

 $r_1 \equiv a_1/a_0$  in  $D^0 \rightarrow K^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-2.40 \pm 0.16$ OUR AVERAGE			
$-2.33 \pm 0.16 \pm 0.08$	71k	1 ABLIKIM	15X BES3 $\ell=e$ , 3-parameter fit
$-2.65 \pm 0.34 \pm 0.08$		BESSION	09 CLEO $\ell=e$ , 3-parameter fit

<sup>1</sup> The 2-parameter fit yields  $-2.23 \pm 0.09 \pm 0.06$ .

 $r_2 \equiv a_2/a_0$  in  $D^0 \rightarrow K^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$5 \pm 4$ OUR AVERAGE			
$3.4 \pm 3.9 \pm 2.4$	71k	ABLIKIM	15X BES3 $\ell=e$ , 3-parameter fit
$13 \pm 9 \pm 1$		BESSION	09 CLEO $\ell=e$ , 3-parameter fit

 $f_+(0)$  in  $D^0 \rightarrow \pi^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.6372 \pm 0.0080 \pm 0.0044$	6.3k	ABLIKIM	15X BES3 $\ell=e$ , 2-parameter fit

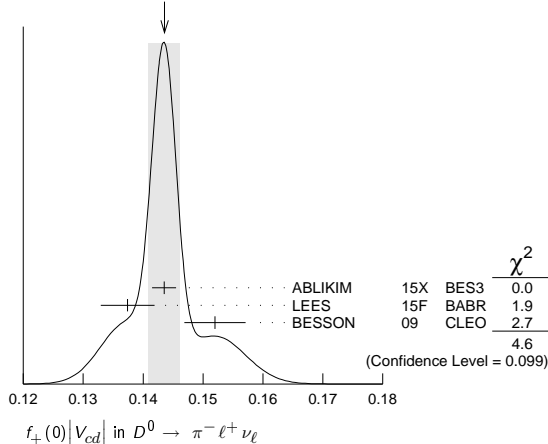
 $f_+(0)|V_{cd}|$  in  $D^0 \rightarrow \pi^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.1436 \pm 0.0026$ OUR AVERAGE			Error includes scale factor of 1.5. See the ideogram below.
$0.1435 \pm 0.0018 \pm 0.0009$	6.3k	1 ABLIKIM	15X BES3 $\ell=e$ , 2-parameter fit
$0.1374 \pm 0.0038 \pm 0.0024$	5.3k	2 LEES	15F BABR $\ell=e$ , 3-parameter fit
$0.152 \pm 0.005 \pm 0.001$		BESSION	09 CLEO $\ell=e$ , 3-parameter fit

<sup>1</sup> The 3-parameter fit yields  $0.1420 \pm 0.0024 \pm 0.0010$ .

<sup>2</sup> LEES 15F reports a value  $0.1374 \pm 0.0038 \pm 0.0022 \pm 0.0009$ , where the last uncertainty is due to the uncertainties of the  $D^0 \rightarrow K^\mp\pi^\pm$  branching fraction.

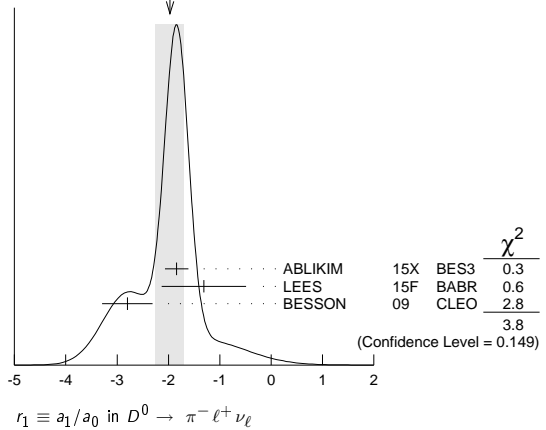
WEIGHTED AVERAGE  
 $0.1436 \pm 0.0026$  (Error scaled by 1.5)

 $r_1 \equiv a_1/a_0$  in  $D^0 \rightarrow \pi^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-1.97 \pm 0.28$ OUR AVERAGE			Error includes scale factor of 1.4. See the ideogram below.
$-1.84 \pm 0.22 \pm 0.07$	6.3k	1 ABLIKIM	15X BES3 $\ell=e$ , 3-parameter fit
$-1.31 \pm 0.70 \pm 0.43$	5.3k	LEES	15F BABR $\ell=e$ , 3-parameter fit
$-2.80 \pm 0.49 \pm 0.04$		BESSION	09 CLEO $\ell=e$ , 3-parameter fit

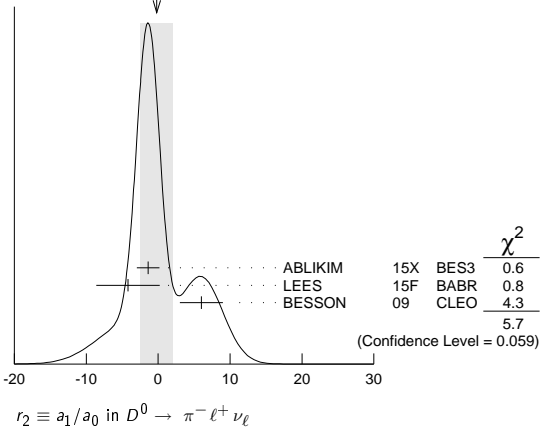
<sup>1</sup> The 2-parameter fit yields  $-2.04 \pm 0.08 \pm 0.03$ .

WEIGHTED AVERAGE  
 $-1.97 \pm 0.28$  (Error scaled by 1.4)

 $r_2 \equiv a_1/a_0$  in  $D^0 \rightarrow \pi^-\ell^+\nu_\ell$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.2 \pm 2.2$ OUR AVERAGE			Error includes scale factor of 1.7. See the ideogram below.
$-1.4 \pm 1.5 \pm 0.5$	6.3k	ABLIKIM	15X BES3 $\ell=e$ , 3-parameter fit
$-4.2 \pm 4.0 \pm 1.9$	5.3k	LEES	15F BABR $\ell=e$ , 3-parameter fit
$6 \pm 3 \pm 0$		BESSION	09 CLEO $\ell=e$ , 3-parameter fit

WEIGHTED AVERAGE  
 $-0.2 \pm 2.2$  (Error scaled by 1.7)

 $D^0$  REFERENCES

STARIC	16	PL B753 412	M. Staric <i>et al.</i>	(BELLE Collab.)
AAIJ	15A	PL B740 158	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15AA	JHEP 1504 043	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15AT	JHEP 1510 055	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABLIKIM	15D	PL B744 339	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15F	PR D91 112015	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15X	PR D92 072012	M. Ablikim <i>et al.</i>	(BES III Collab.)
LEES	15F	PR D91 052022	J.P. Lees <i>et al.</i>	(BABAR Collab.)
MALDE	15	PL B747 9	S. Malde <i>et al.</i>	(BRIS, CERN, MADRA, OXF+)
NAYAK	15	PL B740 1	M. Nayak <i>et al.</i>	(MADRA, OXF, CERN, CMU+)
AAIJ	14AK	JHEP 1407 041	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14AL	PRL 112 041801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14B	PL B728 234	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14BC	JHEP 1410 005	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14C	PL B728 585	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	14Q	PR D90 111103	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABLIKIM	14C	PL B734 227	M. Ablikim <i>et al.</i>	(BES III Collab.)
BONVICINI	14	PR D89 072002	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
KO	14	PRL 112 111801	B.R. Ko <i>et al.</i>	(BELLE Collab.)
LIBBY	14	PL B731 197	J. Libby <i>et al.</i>	(CLEO Collab.)
NISAR	14	PRL 112 211601	N.K. Nisar <i>et al.</i>	(BELLE Collab.)
PENG	14	PR D89 091103	T. Peng <i>et al.</i>	(BELLE Collab.)
TOMARADZE	14	PR D89 031501	A. Tomaradze <i>et al.</i>	(NWES, WAYN)
AAIJ	13AD	PL B723 33	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13AI	PL B725 15	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13BR	PL B726 523	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13CE	PRL 111 251801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13N	PRL 110 101802	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13V	JHEP 1306 065	R. Aaij <i>et al.</i>	(LHCb Collab.)

## Meson Particle Listings

 $D^0$ 

AALTONEN	13AE	PRL 111 231802	T. Aaltonen et al.	(CDF Collab.)	AITALA	01C	PRL 86 3969	E.M. Aitala et al.	(FNAL E791 Collab.)
DOBBS	13	PRL 110 131802	S. Dobbs et al.	(CLEO Collab.)	AITALA	01D	PR D64 112003	E.M. Aitala et al.	(FNAL E791 Collab.)
LEES	13	PR D87 012004	J.P. Lees et al.	(BABAR Collab.)	BONVICINI	01	PR D63 071101	G. Bonvicini et al.	(CLEO Collab.)
LEES	13S	PR D88 071104	J.P. Lees et al.	(BABAR Collab.)	BRANDENB...	01	PRL 87 071802	G. Brandenburg et al.	(CLEO Collab.)
WHITE	13	PR D88 051101	E. White et al.	(BELLE Collab.)	DYTMAN	01	PR D64 111101	S.A. Dytman et al.	(CLEO Collab.)
AAJ	12G	PRL 108 111602	R. Aaij et al.	(LHCb Collab.)	KOPP	01	PR D63 092001	S. Kopp et al.	(CLEO Collab.)
AAJ	12K	JHEP 1204 129	R. Aaij et al.	(LHCb Collab.)	KUSHNIR...	01	PRL 86 5243	A. Kushnirenko et al.	(FNAL SELEX Collab.)
AALTONEN	12AD	PR D86 032007	T. Aaltonen et al.	(CDF Collab.)	LINK	01	PRL 86 2955	J.M. Link et al.	(FNAL FOCUS Collab.)
AALTONEN	12B	PR D85 012009	T. Aaltonen et al.	(CDF Collab.)	BAI	00C	PR D62 052001	J.Z. Bai et al.	(BEP C BES Collab.)
AALTONEN	12O	PRL 109 111801	T. Aaltonen et al.	(CDF Collab.)	GODANG	00	PRL 84 5038	R. Godang et al.	(CLEO Collab.)
ARTUSO	12	PR D85 122002	M. Artuso et al.	(CLEO Collab.)	JUN	00	PRL 84 1857	S.Y. Jun et al.	(FNAL SELEX Collab.)
ASNER	12	PR D86 112001	D.M. Asner, et al	(CLEO Collab.)	LINK	00	PL B485 62	J.M. Link et al.	(FNAL FOCUS Collab.)
INSLER	12	PR D85 092016	J. Insler et al.	(CLEO Collab.)	LINK	00B	PL B491 232	J.M. Link et al.	(FNAL FOCUS Collab.)
LEES	12L	PR D85 091107	J.P. Lees et al.	(BABAR Collab.)	Also		PL B495 443 (errat.)	J.M. Link et al.	(FNAL FOCUS Collab.)
LEES	12Q	PR D86 032001	J.P. Lees et al.	(BABAR Collab.)	PRIPSTEIN	00	PR D61 032005	D. Pripstein et al.	(FNAL E789 Collab.)
PDG	12	PR D86 010001	J. Beringer et al.	(PDG Collab.)	AITALA	99E	PRL 83 32	E.M. Aitala et al.	(FNAL E791 Collab.)
KO	11	PRL 106 211801	B.R. Ko et al.	(BELLE Collab.)	AITALA	99G	PL B462 401	E.M. Aitala et al.	(FNAL E791 Collab.)
LOWREY	11	PR D84 092005	N. Lowrey et al.	(CLEO Collab.)	BONVICINI	99	PRL 82 4586	G. Bonvicini et al.	(CLEO Collab.)
WILLIAMS	11	PR D84 054015	M. Williams	(CLEO Collab.)	AITALA	98	PR D57 13	E.M. Aitala et al.	(FNAL E791 Collab.)
AALTONEN	10X	PR D82 091105	T. Aaltonen et al.	(CDF Collab.)	AITALA	98C	PL B421 405	E.M. Aitala et al.	(FNAL E791 Collab.)
ANASHIN	10A	PL B686 84	V.V. Anashin et al.	(VEPP-4M KEDR Collab.)	AITALA	98D	PL B423 185	E.M. Aitala et al.	(FNAL E791 Collab.)
ASNER	10	PR D81 052007	D.M. Asner et al.	(CLEO Collab.)	ARTUSO	98	PRL 80 3193	M. Artuso et al.	(CLEO Collab.)
BHATTACHAR.	10A	PR D81 096008	B. Bhattacharya, C.-W. Chiang, J.L. Rosner	(CHIC+ Collab.)	ASNER	98	PR D58 092001	D.M. Asner et al.	(CLEO Collab.)
DEL-AMO-SA...	10	PR D81 111103	P. del Amo Sanchez et al.	(BABAR Collab.)	BARATE	98W	PL B436 211	R. Barate et al.	(ALEPH Collab.)
DEL-AMO-SA...	10D	PRL 105 081803	P. del Amo Sanchez et al.	(BABAR Collab.)	COAN	98	PRL 80 1150	T.E. Coan et al.	(CLEO Collab.)
MENDEZ	10	PR D81 052013	H. Mendez et al.	(CLEO Collab.)	PDG	98	EPJ C 3 1	C. Caso et al.	(PDG Collab.)
PETRIC	10	PR D81 091102	M. Petric et al.	(BELLE Collab.)	ADAMOVICH	97	PL B408 469	M.I. Adamovich et al.	(CERN BEATRICE Collab.)
AUBERT	09AI	PR D80 071103	B. Aubert et al.	(BABAR Collab.)	BARATE	97C	PL B403 367	R. Barate et al.	(ALEPH Collab.)
AUBERT	09AM	PRL 103 211801	B. Aubert et al.	(BABAR Collab.)	AITALA	96C	PRL 77 2384	E.M. Aitala et al.	(FNAL E791 Collab.)
BEDIAGA	09	PR D80 096006	I. Bediaga et al.	(CBP NDAM Collab.)	ALBRECHT	96C	PL B374 249	H. Albrecht et al.	(ARGUS Collab.)
BESSON	09	PR D80 032005	D. Besson et al.	(CLEO Collab.)	ALEXOPOU...	96	PRL 77 2380	T. Alexopoulos et al.	(FNAL E771 Collab.)
Also		PR D79 052010	J.Y. Ge et al.	(CLEO Collab.)	ASNER	96B	PR D54 4211	D.M. Asner et al.	(CLEO Collab.)
LOWREY	09	PR D80 031105	N. Lowrey et al.	(CLEO Collab.)	BARISH	96	PL B373 334	B.C. Barish et al.	(CLEO Collab.)
RUBIN	09	PR D79 097101	P. Rubin et al.	(CLEO Collab.)	FRABETTI	96B	PL B382 312	P.L. Frabetti et al.	(FNAL E687 Collab.)
ZUPANEC	09	PR D80 052006	A. Zupanc et al.	(BELLE Collab.)	Also		PRL 77 2147 (erratum)	A. Freyberger et al.	(CLEO Collab.)
AALTONEN	08E	PRL 100 121802	T. Aaltonen et al.	(CDF Collab.)	FREYBERGER	96	PRL 76 3065	A. Freyberger et al.	(CLEO Collab.)
ABLIKIM	08L	PL B665 16	M. Ablikim et al.	(BES Collab.)	KUBOTA	96B	PR D54 2994	Y. Kubota et al.	(CLEO Collab.)
ARINSTEIN	08	PL B662 102	K. Arinstein et al.	(BELLE Collab.)	ADAMOVICH	95	PL B353 563	M.I. Adamovich et al.	(CERN BEATRICE Collab.)
ARTUSO	08	PR D77 092003	M. Artuso et al.	(CLEO Collab.)	BARTELT	95	PR D52 4860	J.E. Bartelt et al.	(CLEO Collab.)
ASNER	08	PR D78 012001	D.M. Asner et al.	(CLEO Collab.)	BUTLER	95	PR D52 2656	F. Butler et al.	(CLEO Collab.)
AUBERT	08AL	PR D78 034023	B. Aubert et al.	(BABAR Collab.)	FRABETTI	95C	PL B354 486	P.L. Frabetti et al.	(FNAL E687 Collab.)
AUBERT	08AO	PR D78 051102	B. Aubert et al.	(BABAR Collab.)	FRABETTI	95G	PL B364 127	P.L. Frabetti et al.	(FNAL E687 Collab.)
AUBERT	08AZ	PR D78 071101	B. Aubert et al.	(BABAR Collab.)	KODAMA	95	PL B345 85	K. Kodama et al.	(FNAL E653 Collab.)
AUBERT	08L	PRL 100 051802	B. Aubert et al.	(BABAR Collab.)	ALBRECHT	94	PL B324 249	H. Albrecht et al.	(ARGUS Collab.)
AUBERT	08M	PRL 100 061803	B. Aubert et al.	(BABAR Collab.)	ALBRECHT	94F	PL B340 125	H. Albrecht et al.	(ARGUS Collab.)
AUBERT	08U	PR D78 011105	B. Aubert et al.	(BABAR Collab.)	ALBRECHT	94I	ZPHY C64 375	H. Albrecht et al.	(ARGUS Collab.)
BITENC	08	PR D77 112003	U. Bitenc et al.	(BELLE Collab.)	FRABETTI	94C	PL B321 295	P.L. Frabetti et al.	(FNAL E687 Collab.)
BONVICINI	08	PR D77 091106	G. Bonvicini et al.	(CLEO Collab.)	FRABETTI	94D	PL B323 459	P.L. Frabetti et al.	(FNAL E687 Collab.)
DOBBS	08	PR D77 112005	S. Dobbs et al.	(CLEO Collab.)	FRABETTI	94G	PL B331 217	P.L. Frabetti et al.	(FNAL E687 Collab.)
Also		PRL 100 251802	D. Cronin-Hennessy et al.	(CLEO Collab.)	FRABETTI	94J	PL B340 254	P.L. Frabetti et al.	(FNAL E687 Collab.)
GASPERO	08	PR D78 014015	M. Gaspero et al.	(ROMA, CINN, TELA Collab.)	KODAMA	94	PL B336 605	K. Kodama et al.	(FNAL E653 Collab.)
HE	08	PRL 100 091801	Q. He et al.	(CLEO Collab.)	MISHRA	94	PR D50 89	C.S. Mishra et al.	(FNAL E789 Collab.)
PDG	08	PL B667 1	C. Amers et al.	(PDG Collab.)	AKERIB	93C	PR D71 3070	D.S. Akerib et al.	(CLEO Collab.)
STARIC	08	PL B670 190	M. Staric et al.	(BELLE Collab.)	ALBRECHT	93D	PL B308 435	H. Albrecht et al.	(ARGUS Collab.)
ABLIKIM	07G	PL B658 1	M. Ablikim et al.	(BES Collab.)	ANJOS	93	PR D48 56	J.C. Anjos et al.	(FNAL E691 Collab.)
ARTUSO	07A	PRL 99 191801	M. Artuso et al.	(CLEO Collab.)	BEAN	93C	PL B317 647	A. Bean et al.	(CLEO Collab.)
AUBERT	07AB	PR D76 014018	B. Aubert et al.	(BABAR Collab.)	FRABETTI	93I	PL B315 203	P.L. Frabetti et al.	(FNAL E687 Collab.)
AUBERT	07BG	PR D76 052005	B. Aubert et al.	(BABAR Collab.)	KODAMA	93B	PL B313 260	K. Kodama et al.	(FNAL E653 Collab.)
AUBERT	07BJ	PRL 99 251801	B. Aubert et al.	(BABAR Collab.)	PROCARIO	93B	PR D48 4007	M. Procaro et al.	(CLEO Collab.)
AUBERT	07T	PR D76 011102	B. Aubert et al.	(BABAR Collab.)	SELEN	93	PRL 71 1973	M.A. Selen et al.	(CLEO Collab.)
AUBERT	07W	PRL 98 211802	B. Aubert et al.	(BABAR Collab.)	ADAMOVICH	92	PL B280 163	M.I. Adamovich et al.	(CERN WA82 Collab.)
CRAWFIELD	07	PRL 98 092002	C. Crawford et al.	(CLEO Collab.)	ALBRECHT	92P	ZPHY C56 7	H. Albrecht et al.	(ARGUS Collab.)
DOBBS	07	PR D76 112001	S. Dobbs et al.	(CLEO Collab.)	ANJOS	92B	PR D46 R1	J.C. Anjos et al.	(FNAL E691 Collab.)
LINK	07A	PR D75 052013	J.M. Link et al.	(FNAL FOCUS Collab.)	ANJOS	92C	PR D46 1941	J.C. Anjos et al.	(FNAL E691 Collab.)
STARIC	07	PRL 98 211803	M. Staric et al.	(BELLE Collab.)	BARLAG	92C	ZPHY C55 383	S. Barlag et al.	(ACCMOR Collab.)
ZHANG	07B	PR 99 131803	L.M. Zhang et al.	(BELLE Collab.)	Also		ZPHY C48 29	S. Barlag et al.	(ACCMOR Collab.)
ABLIKIM	06O	EPJ C47 31	M. Ablikim et al.	(BES Collab.)	COFFMAN	92B	PR D45 2196	D.M. Coffman et al.	(Mark III Collab.)
ABLIKIM	06U	PL B643 246	M. Ablikim et al.	(BES Collab.)	Also		PRL 64 2615	J. Adler et al.	(Mark III Collab.)
ABULENCIA	06X	PR D74 031109	A. Abulencia et al.	(CDF Collab.)	FRABETTI	92	PL B281 167	P.L. Frabetti et al.	(FNAL E687 Collab.)
ADAM	06A	PRL 97 251801	N.E. Adam et al.	(CLEO Collab.)	FRABETTI	92B	PL B286 195	P.L. Frabetti et al.	(FNAL E687 Collab.)
AUBERT.B	06N	PRL 97 221803	B. Aubert et al.	(BABAR Collab.)	ALVAREZ	91B	ZPHY C50 11	M.P. Alvarez et al.	(CERN NA14/2 Collab.)
AUBERT.B	06X	PR D74 091102	B. Aubert et al.	(BABAR Collab.)	AMMAR	91	PR D44 3383	R. Ammar et al.	(CLEO Collab.)
CRAWFIELD	06A	PR D74 031108	C. Crawford et al.	(CLEO Collab.)	ANJOS	91	PR D43 R635	J.C. Anjos et al.	(FNAL-TPS Collab.)
HUANG	06B	PR D74 112005	G.S. Huang et al.	(CLEO Collab.)	ANJOS	91D	PR D44 R3371	J.C. Anjos et al.	(FNAL-TPS Collab.)
PDG	06	IP 633 1	W.-M. Yao et al.	(PDG Collab.)	BARLAG	91	PR D46 1011	Z. Barlag et al.	(Mark III Collab.)
RUBIN	06	PRL 96 081802	P. Rubin et al.	(CLEO Collab.)	COFFMAN	91	PL B263 135	D.M. Coffman et al.	(Mark III Collab.)
WIDHALM	06	PRL 97 061804	L. Widhalm et al.	(BELLE Collab.)	CRAWFORD	91B	PR D44 3394	G. Crawford et al.	(CLEO Collab.)
ZHANG	06	PRL 96 151801	L.M. Zhang et al.	(BELLE Collab.)	DECAMP	91J	PL B266 218	D. Decamp et al.	(ALEPH Collab.)
ABLIKIM	05F	PL B622 6	M. Ablikim et al.	(BES Collab.)	FRABETTI	91	PL B263 584	P.L. Frabetti et al.	(FNAL E687 Collab.)
ABLIKIM	05P	PL B625 196	M. Ablikim et al.	(BES Collab.)	KINOSHITA	91	PR D43 2836	K. Kinoshita et al.	(CLEO Collab.)
ACOSTA	05C	PRL 94 122001	D. Acosta et al.	(FNAL CDF Collab.)	KODAMA	91	PRL 66 1819	K. Kodama et al.	(FNAL E653 Collab.)
ASNER	05	PR D72 012001	D.M. Asner et al.	(CLEO Collab.)	ALBRECHT	90C	ZPHY C46 9	H. Albrecht et al.	(ARGUS Collab.)
AUBERT.B	05J	PR D72 052008	B. Aubert et al.	(BABAR Collab.)	ALEXANDER	90	PRL 65 1184	J. Alexander et al.	(CLEO Collab.)
BITEN C	05	PR D72 071101	U. Bitenc et al.	(BELLE Collab.)	ALVAREZ	90	ZPHY C47 539	M.P. Alvarez et al.	(CERN NA14/2 Collab.)
CRAWFIELD	05	PR D71 077101	C. Crawford et al.	(CLEO Collab.)	ANJOS	90D	PR D42 2414	J.C. Anjos et al.	(FNAL E691 Collab.)
COAN	05	PRL 95 181802	T.E. Coan et al.	(CLEO Collab.)	BARLAG	90C	ZPHY C46 563	S. Barlag et al.	(ACCMOR Collab.)
CRONIN-HEN...	05	PR D72 031102	D. Cronin-Hennessy et al.	(CLEO Collab.)	ADLER	89	PRL 62 1821	J. Adler et al.	(Mark III Collab.)
HE	05	PRL 95 121801	Q. He et al.	(CLEO Collab.)	ADLER	89C	PR D40 906	J. Adler et al.	(Mark III Collab.)
Also		PRL 96 199903 (errat.)	Q. He et al.	(CLEO Collab.)	ALBRECHT	89D	ZPHY C43 181	H. Albrecht et al.	(ARGUS Collab.)
HUANG	05	PRL 94 011802	G.S. Huang et al.	(CLEO Collab.)	ANJOS	89F	PRL 62 1587	J.C. Anjos et al.	(FNAL E691 Collab.)
KAYIS-TOPAK...	05	PL B626 24	A. Kayis-Topaksu et al.	(CERN CHORUS Collab.)	ABACHI	88	PL B205 411	S. Abachi et al.	(HRS Collab.)
LI	05A	PRL 94 071801	J. Li et al.	(BELLE Collab.)	ADLER	88	PR D37 2023	J. Adler et al.	(Mark III Collab.)
LINK	05	PL B607 51	J.M. Link et al.	(FNAL FOCUS Collab.)	ADLER	88C	PRL 60 89	J. Adler et al.	(Mark III Collab.)
LINK	05A	PL B607 59	J.M. Link et al.	(FNAL FOCUS Collab.)	ALBRECHT	88G	PL B209 380	H. Albrecht et al.	(ARGUS Collab.)
LINK	05B	PL B607 67	J.M. Link et al.	(FNAL FOCUS Collab.)	ALBRECHT	88I	PL B210 267	H. Albrecht et al.	(ARGUS Collab.)
LINK	05E	PL B622 239	J.M. Link et al.	(FNAL FOCUS Collab.)	ANJOS	88C	PRL 60 1239	J.C. Anjos et al.	(FNAL E691 Collab.)
LINK	05L	PL B610 225	J.M. Link et al.	(FNAL FOCUS Collab.)	BORTOLETTO	88	PR D37 1719	D. Bortoletto et al.	(CLEO Collab.)
LINK	05H	PL B618 23	J.M. Link et al.	(FNAL FOCUS Collab.)	Also		PR D39 1471 (erratum)	D. Bortoletto et al.	(CLEO Collab.)
ONENGUT	05	PL B613 105	G. Onengut et al.	(CERN CHORUS Collab.)	HAAS	88	PRL 60 1614	P. Haas et al.	(CLEO Collab.)
TIAN	05	PRL 95 231801	X.C. Tian et al.	(BELLE Collab.)	RAAB	88	PR D37 2391	J.R. Raab et al.	(FNAL E691 Collab.)
ABLIKIM	04C	PL B597 39	M. Ablikim et al.	(BEP C BES Collab.)	ADLER	87	PL B196 107	J. Adler et al.	(Mark III Collab.)
ABT	04	PL B596 173	I. Abt et al.	(HERA B Collab.)	AGUILAR-...	87E	ZPHY C36 551	M. Aguilar-Benitez et al.	(LEBC-EHS Collab.)
ASNER	04A	PR D70 091101	D.M. Asner et al.	(CLEO Collab.)	Also		ZPHY C40 321	M. Aguilar-Benitez et al.	(LEBC-EHS Collab.)
AUBERT	04Q	PR D69 051101	B. Aubert et al.	(BABAR Collab.)	AGUILAR-...	87F	ZPHY C36 559	M. Aguilar-Benitez et al.	(LEBC-EHS Collab.)
AUBERT.B	04Q	PR D70 091102	B. Aubert et al.	(BABAR Collab.)	Also		ZPHY C38 520 (erratum)	M. Aguilar-Benitez et al.	(LEBC-EHS Collab.)
AUBERT.B	04Y	PRL 93 191801	B. Aubert et al.	(BABAR Collab.)	BARLAG	87B	ZPHY C37 17	S. Barlag et al.	(ACCMOR Collab.)
LINK	04B	PL B586 21	J.M. Link et al.	(FNAL FOCUS Collab.)	BECKER	87C	PL B193 147	J.J. Becker et al.	(Mark III Collab.)
LINK	04D	PL B586 191	J.M. Link et al.	(FNAL FOCUS Collab.)	LINK	87C	PL B198 590 (erratum)	J.J. Becker et al.	(Mark III Collab.)
RUBIN	04	PRL 93 111801	P. Rubin et al.	(CLEO Collab.)	PAKA	87	PL B189 238	H. Paka et al.	(ACCMOR Collab.)
TAJIMA	04	PRL 92 101803	O. Tajima et al.	(BELLE Collab.)	RILES	87	PR D35 2914	K. Riles et al.	(Mark II Collab.)
ACOSTA	03F	PR D68 091101	D. Acosta et al.	(CDF Collab.)	BAILEY	86	ZPHY C30 51	R. Bailey et al.	(CLEO Collab.)
AUBERT	03P	PRL 91 121801	B. Aubert et al.	(BABAR Collab.)	BEBEK	86	PRL 56 1893	C. Bebek et al.	(CLEO Collab.)
AUBERT	03Z	PRL 91 171801	B. Aubert et al.	(BABAR Collab.)	LOUIS	86	PRL 56 1027	W.C. Louis et al.	(PRIN, CHIC, ISU Collab.)
COAN									

See key on page 601

## Meson Particle Listings

$$D^0, D^*(2007)^0, D^*(2010)^\pm$$

AVERY	80	PRL 44 1309	P. Avery <i>et al.</i>	(ILL, FNAL, COLU)
ABRAMS	79D	PRL 43 481	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
VUILLEMIN	78	PRL 41 1149	V. Vuillemin <i>et al.</i>	(LGV Collab.)
PERUZZI	77	PRL 39 1301	I. Peruzzi <i>et al.</i>	(LGV Collab.)
PICCOLO	77	PL 70B 260	M. Piccolo <i>et al.</i>	(Mark I Collab.)
MOSTELLER	48	Am.Stat.3 No.5 30	R.A. Fisher, F. Mosteller	

## OTHER RELATED PAPERS

RICHMAN	95	RMP 67 893	J.D. Richman, P.R. Burchat	(UCSB, STAN)
ROSNER	95	CNPP 21 369	J. Rosner	(CHIC)

$$D^*(2007)^0 \quad I(J^P) = \frac{1}{2}(1^-)$$

$I, J, P$  need confirmation.

$J$  consistent with 1, value 0 ruled out (NGUYEN 77).

 $D^*(2007)^0$  MASS

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2006.85 ± 0.05 OUR FIT</b>				Error includes scale factor of 1.1.
• • •				We do not use the following data for averages, fits, limits, etc. • • •
2006 ± 1.5		<sup>1</sup> GOLDHABER 77	MRK1	$e^+e^-$
		<sup>1</sup> From simultaneous fit to $D^*(2010)^+, D^*(2007)^0, D^+,$ and $D^0$ .		

 $m_{D^*(2007)^0} - m_{D^0}$ 

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>142.016 ± 0.030 OUR FIT</b>				Error includes scale factor of 1.5.
<b>142.016 ± 0.030 OUR AVERAGE</b>				Error includes scale factor of 1.5.
142.007 ± 0.015 ± 0.014	10K	<sup>2</sup> TOMARADZE 15	CLEO	$e^+e^- \rightarrow$ hadrons
142.2 ± 0.3 ± 0.2	145	ALBRECHT 95F	ARG	$e^+e^- \rightarrow$ hadrons
142.12 ± 0.05 ± 0.05	1176	BORTOLETTO 92B	CLE2	$e^+e^- \rightarrow$ hadrons
• • •				We do not use the following data for averages, fits, limits, etc. • • •
142.2 ± 2.0		SADROZINSKI 80	CBAL	$D^{*0} \rightarrow D^0\pi^0$
142.7 ± 1.7		<sup>3</sup> GOLDHABER 77	MRK1	$e^+e^-$

<sup>2</sup> Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration. This value comes from the average of the results for two decay modes,  $D^0 \rightarrow K^-\pi^+\pi^+$  and  $D^0 \rightarrow K^-\pi^+\pi^-\pi^+$ .

<sup>3</sup> From simultaneous fit to  $D^*(2010)^+, D^*(2007)^0, D^+,$  and  $D^0$ .

 $D^*(2007)^0$  WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.1</b>	90	<sup>4</sup> ABACHI 88B	HRS	$D^{*0} \rightarrow D^+\pi^-$

<sup>4</sup> Assuming  $m_{D^{*0}} = 2007.2 \pm 2.1$  MeV/ $c^2$ .

 $D^*(2007)^0$  DECAY MODES

$\bar{D}^*(2007)^0$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad D^0\pi^0$	(64.7 ± 0.9) %
$\Gamma_2 \quad D^0\gamma$	(35.3 ± 0.9) %

## CONSTRAINED FIT INFORMATION

An overall fit to 2 branching ratios uses 5 measurements and one constraint to determine 2 parameters. The overall fit has a  $\chi^2 = 2.5$  for 4 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$$\begin{array}{c|c} x_2 & -100 \\ \hline & x_1 \end{array}$$

 $D^*(2007)^0$  BRANCHING RATIOS

$\Gamma(D^0\pi^0)/\Gamma(D^0\gamma)$					$\Gamma_1/\Gamma_2$
VALUE		EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.83 ± 0.07 OUR FIT</b>					Error includes scale factor of 1.1.
<b>1.85 ± 0.07 OUR AVERAGE</b>					
1.90 ± 0.07 ± 0.05	4.9k		ABLIKIM 15B	BES3	10.6 $e^+e^- \rightarrow$ hadrons
1.74 ± 0.02 ± 0.13			AUBERT_BE 05G	BABR	10.6 $e^+e^- \rightarrow$ hadrons

$\Gamma(D^0\pi^0)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.647±0.009 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.655 ± 0.008 ± 0.005	3.2k	<sup>5</sup> ABLIKIM	15B	BES3 $e^+e^- \rightarrow$ hadrons
0.635 ± 0.003 ± 0.017	69k	<sup>5</sup> AUBERT, BE	05G	BABR 10.6 $e^+e^- \rightarrow$ hadrons
0.596 ± 0.035 ± 0.028	858	<sup>6</sup> ALBRECHT	95F	ARG $e^+e^- \rightarrow$ hadrons
0.636 ± 0.023 ± 0.033	1097	<sup>6</sup> BUTLER	92	CLE2 $e^+e^- \rightarrow$ hadrons

$\Gamma(D^0\gamma)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.353 ± 0.009 OUR FIT</b>				
<b>0.381 ± 0.029 OUR AVERAGE</b>				
0.404 ± 0.035 ± 0.028	456	<sup>6</sup> ALBRECHT	95F ARG	$e^+e^- \rightarrow$ hadrons
0.364 ± 0.023 ± 0.033	621	<sup>6</sup> BUTLER	92 CLE2	$e^+e^- \rightarrow$ hadrons
0.37 ± 0.08 ± 0.08		ADLER	88D MRK3	$e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.345 ± 0.008 ± 0.005	1.8k	<sup>5</sup> ABLIKIM	15B BES3	$e^+e^- \rightarrow$ hadrons
0.365 ± 0.003 ± 0.017	68k	<sup>5</sup> AUBERT,BE	05G BABR	10.6 $e^+e^- \rightarrow$ hadrons
0.47 ± 0.23		LOW	87 HRS	29 GeV $e^+e^-$
0.53 ± 0.13		BARTEL	85G JADE	$e^+e^-$ , hadrons
0.47 ± 0.12		COLES	82 MRK2	$e^+e^-$
0.45 ± 0.15		GOLDHABER	77 MRK1	$e^+e^-$

<sup>5</sup> Derived from the ratio  $\Gamma(D^0\pi^0)/\Gamma(D^0\gamma)$  assuming that the branching fractions of  $D^{*0} \rightarrow D^0\pi^0$  and  $D^{*0} \rightarrow D^0\gamma$  decays sum to 100%

<sup>6</sup> The BUTLER 92 and ALBRECHT 95F branching ratios are not independent, they have been constrained by the authors to sum to 100%.

 $D^*(2007)^0$  REFERENCES

ABLIKIM 15B	PR D91 031101	M. Ablikim <i>et al.</i>	(BES III Collab.)
TOMARADZE 15	PR D91 011102	A. Tomaradze <i>et al.</i>	(NWES)
AUBERT, BE 05G	PR D72 091101	B. Aubert <i>et al.</i>	(BABAR Collab.)
ALBRECHT 95F	ZPHY C66 63	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BORTOLETTO 92B	PRL 69 2046	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
BUTLER 92	PRL 69 2041	F. Butler <i>et al.</i>	(CLEO Collab.)
ABACHI 88B	PL B212 533	S. Abachi <i>et al.</i>	(ANL, IND, MICH, PURD+)
ADLER 88D	PL B208 152	J. Adler <i>et al.</i>	(Mark III Collab.)
LOW 87	PL B183 232	E.H. Low <i>et al.</i>	(HRS Collab.)
BARTEL 85G	PL 161B 197	W. Bartel <i>et al.</i>	(JADE Collab.)
COLES 82	PR D26 2190	M.W. Coles <i>et al.</i>	(LBL, SLAC)
SADROZINSKI 80	Madison Conf. 681	H.F.W. Sadrozinski <i>et al.</i>	(PRIN, CIT+)
GOLDHABER 77	PL 69B 503	G. Goldhaber <i>et al.</i>	(Mark I Collab.)
NGUYEN 77	PRL 39 262	H.K. Nguyen <i>et al.</i>	(LBL, SLAC) J

$$D^*(2010)^\pm \quad I(J^P) = \frac{1}{2}(1^-)$$

$I, J, P$  need confirmation.

 $D^*(2010)^\pm$  MASS

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>2010.26 ± 0.05 OUR FIT</b>				
• • •				We do not use the following data for averages, fits, limits, etc. • • •
2008 ± 3	<sup>1</sup> GOLDHABER 77	MRK1	±	$e^+e^-$
2008.6 ± 1.0	<sup>2</sup> PERUZZI 77	LGW	±	$e^+e^-$
	<sup>1</sup> From simultaneous fit to $D^*(2010)^+, D^*(2007)^0, D^+,$ and $D^0$ ; not independent of FELDMAN 77B mass difference below.			
	<sup>2</sup> PERUZZI 77 mass not independent of FELDMAN 77B mass difference below and PERUZZI 77 $D^0$ mass value.			

 $m_{D^*(2010)^+} - m_{D^+}$ 

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>140.68 ± 0.08 OUR FIT</b>				
<b>140.64 ± 0.08 ± 0.06</b>	620	BORTOLETTO 92B	CLE2	$e^+e^- \rightarrow$ hadrons

 $m_{D^*(2010)^+} - m_{D^0}$ 

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>145.4257 ± 0.0017 OUR FIT</b>				
<b>145.4258 ± 0.0020 OUR AVERAGE</b>				Error includes scale factor of 1.2.
145.4259 ± 0.0004 ± 0.0017	312.8k	LEES 13x	BABR	$D^{*\pm} \rightarrow D^0\pi^\pm \rightarrow (K\pi, K3\pi)\pi^\pm$
145.412 ± 0.002 ± 0.012		ANASTASSOV 02	CLE2	$D^{*\pm} \rightarrow D^0\pi^\pm \rightarrow (K\pi)\pi^\pm$
145.54 ± 0.08	611	<sup>3</sup> ADINOLFI 99	BEAT	$D^{*\pm} \rightarrow D^0\pi^\pm$
145.45 ± 0.02		<sup>3</sup> BREITWEG 99	ZEUS	$D^{*\pm} \rightarrow D^0\pi^\pm \rightarrow (K\pi)\pi^\pm$

# Meson Particle Listings

## $D^*(2010)^\pm, D_0^*(2400)^0$

145.42 ±0.05		3	BREITWEG	99	ZEUS	$D^{*\pm} \rightarrow D^0 \pi^\pm \rightarrow (K^- 3\pi) \pi^\pm$
145.5 ±0.15	103	4	ADLOFF	97B	H1	$D^{*\pm} \rightarrow D^0 \pi^\pm$
145.44 ±0.08	152	4	BREITWEG	97	ZEUS	$D^{*\pm} \rightarrow D^0 \pi^\pm, D^{*0} \rightarrow K^- 3\pi, D^{*+} \rightarrow D^0 \pi^+, D^{*0} \rightarrow K^- \pi^+$
145.42 ±0.11	199	4	BREITWEG	97	ZEUS	$D^{*+} \rightarrow D^0 \pi^+, D^{*0} \rightarrow K^- \pi^+$
145.4 ±0.2	48	4	DERRICK	95	ZEUS	$D^{*+} \rightarrow D^0 \pi^+$
145.39 ±0.06 ±0.03			BARLAG	92B	ACCM	$\pi^- 230 \text{ GeV}$
145.5 ±0.2	115	4	ALEXANDER	91B	OPAL	$D^{*+} \rightarrow D^0 \pi^+$
145.30 ±0.06		4	DECAMP	91J	ALEP	$D^{*+} \rightarrow D^0 \pi^+$
145.40 ±0.05 ±0.10			ABACHI	88B	HRS	$D^{*+} \rightarrow D^0 \pi^+$
145.46 ±0.07 ±0.03			ALBRECHT	85F	ARG	$D^{*+} \rightarrow D^0 \pi^+$
145.5 ±0.3	28		BAILEY	83	SPEC	$D^{*+} \rightarrow D^0 \pi^+$
145.5 ±0.3	60		FITCH	81	SPEC	$\pi^- A$
145.3 ±0.5	30		FELDMAN	77B	MRK1	$D^{*+} \rightarrow D^0 \pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
145.4256±0.0006±0.0017 138.5k			LEES	13x	BABR	$D^{*\pm} \rightarrow D^0 \pi^\pm \rightarrow (K^- \pi^+) \pi^\pm$
145.4266±0.0005±0.0019 174.3k			LEES	13x	BABR	$D^{*\pm} \rightarrow D^0 \pi^\pm \rightarrow (K^- 2\pi^+ \pi^-) \pi^\pm$
145.44 ±0.09	122	4	BREITWEG	97B	ZEUS	$D^{*+} \rightarrow D^0 \pi^+, D^{*0} \rightarrow K^- \pi^+$
145.8 ±1.5	16		AHLEN	83	HRS	$D^{*+} \rightarrow D^0 \pi^+$
145.1 ±1.8	12		BAILEY	83	SPEC	$D^{*+} \rightarrow D^0 \pi^+$
145.1 ±0.5	14		BAILEY	83	SPEC	$D^{*+} \rightarrow D^0 \pi^+$
145.5 ±0.5	14		YELTON	82	MRK2	$29 e^+ e^- \rightarrow K^- \pi^+$
~ 145.5			AVERY	80	SPEC	$\gamma A$
145.2 ±0.6	2		BLIETSCHAU	79	BEBC	$\nu p$
3 Statistical errors only.						
4 Systematic error not evaluated.						

### $m_{D^*(2010)^+} - m_{D^*(2007)^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.6±1.8	5	PERUZZI 77	LGW $e^+ e^-$
5 Not independent of FELDMAN 77B mass difference above, PERUZZI 77 $D^0$ mass, and GOLDHABER 77 $D^*(2007)^0$ mass.			

### $D^*(2010)^\pm$ WIDTH

VALUE (keV)	CL% EVTS	DOCUMENT ID	TECN	COMMENT
<b>83.4±1.8 OUR AVERAGE</b>				
83.3±1.2± 1.4	312.8k	6	LEES	13x BABR $D^{*\pm} \rightarrow D^0 \pi^\pm \rightarrow (K \pi, K 3\pi) \pi^\pm$
96 ±4 ±22		6	ANASTASSOV 02	CLE2 $D^{*\pm} \rightarrow D^0 \pi^\pm \rightarrow (K \pi) \pi^\pm$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
83.4±1.7± 1.5	138.5k	6	LEES	13x BABR $D^{*+} \rightarrow D^0 \pi^+ \rightarrow (K^- \pi^+) \pi^+$
83.2±1.5± 2.6	174.3k	6	LEES	13x BABR $D^{*+} \rightarrow D^0 \pi^+ \rightarrow (K^- 2\pi^+ \pi^-) \pi^+$
<131	90 110	BARLAG	92B ACCM	$\pi^- 230 \text{ GeV}$
6 Ignoring the electromagnetic contribution from $D^{*+} \rightarrow D^\pm \gamma$ .				

### $D^*(2010)^\pm$ DECAY MODES

$D^*(2010)^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 D^0 \pi^+$	(67.7±0.5) %
$\Gamma_2 D^+ \pi^0$	(30.7±0.5) %
$\Gamma_3 D^+ \gamma$	( 1.6±0.4) %

### CONSTRAINED FIT INFORMATION

An overall fit to 3 branching ratios uses 6 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 0.3$  for 4 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−62	
$x_3$	−43	−44
	$x_1$	$x_2$

### $D^*(2010)^+$ BRANCHING RATIOS

$\Gamma(D^0 \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE				
<b>0.677 ±0.005 OUR FIT</b>				
<b>0.677 ±0.006 OUR AVERAGE</b>				
0.6759±0.0029±0.0064	7,8,9	BARTELT	98 CLE2	$e^+ e^-$
0.688 ±0.024 ±0.013		ALBRECHT	95F ARG	$e^+ e^- \rightarrow$ hadrons
0.681 ±0.010 ±0.013	7	BUTLER	92 CLE2	$e^+ e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.57 ±0.04 ±0.04		ADLER	88D MRK3	$e^+ e^-$
0.44 ±0.10		COLES	82 MRK2	$e^+ e^-$
0.6 ±0.15	9	GOLDHABER	77 MRK1	$e^+ e^-$

$\Gamma(D^+ \pi^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE				
<b>0.307 ±0.005 OUR FIT</b>				
<b>0.3073±0.0013±0.0062</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.312 ±0.011 ±0.008	1404	ALBRECHT	95F ARG	$e^+ e^- \rightarrow$ hadrons
0.308 ±0.004 ±0.008	410	7 BUTLER	92 CLE2	$e^+ e^- \rightarrow$ hadrons
0.26 ±0.02 ±0.02		ADLER	88D MRK3	$e^+ e^-$
0.34 ±0.07		COLES	82 MRK2	$e^+ e^-$

$\Gamma(D^+ \gamma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
VALUE	CL% EVTS			
<b>0.016 ±0.004 OUR FIT</b>				
<b>0.016 ±0.005 OUR AVERAGE</b>				
0.0168±0.0042±0.0029		7,8	BARTELT	98 CLE2 $e^+ e^-$
0.011 ±0.014 ±0.016	12	7	BUTLER	92 CLE2 $e^+ e^- \rightarrow$ hadrons

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.052 90 ALBRECHT 95F ARG  $e^+ e^- \rightarrow$  hadrons

0.17 ±0.05 ±0.05 ADLER 88D MRK3  $e^+ e^-$

0.22 ±0.12 10 COLES 82 MRK2  $e^+ e^-$

7 The branching ratios are not independent, they have been constrained by the authors to sum to 100%.

8 Systematic error includes theoretical error on the prediction of the ratio of hadronic modes.

9 Assuming that isospin is conserved in the decay.

10 Not independent of  $\Gamma(D^0 \pi^+)/\Gamma_{\text{total}}$  and  $\Gamma(D^+ \pi^0)/\Gamma_{\text{total}}$  measurement.

### $D^*(2010)^\pm$ REFERENCES

LEES	13X	PRL 111 111801	J.P. Lees <i>et al.</i>	(BABAR Collab.)
Also		PR D88 052003	J.P. Lees <i>et al.</i>	(BABAR Collab.)
Also		PR D88 079902 (err.)	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ANASTASSOV	02	NP B547 3	A. Anastassov <i>et al.</i>	(CLEO Collab.)
ADINOLFI	99	EPJ C6 67	M. Adinolfi <i>et al.</i>	(Beatrice Collab.)
BREITWEG	99	PRL 80 3919	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
BARTELT	98	ZPHY C72 593	C. Adloff <i>et al.</i>	(H1 Collab.)
ADLOFF	97B	PL B401 192	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
BREITWEG	97B	PL B407 402	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
ALBRECHT	95F	ZPHY C66 63	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DERRICK	95	PL B349 225	M. Derrick <i>et al.</i>	(ZEUS Collab.)
BARLAG	92B	PL B278 480	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
BORTOLETTO	92B	PRL 69 2046	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
BUTLER	92	PRL 69 2041	F. Butler <i>et al.</i>	(CLEO Collab.)
ALEXANDER	91B	PL B262 341	G. Alexander <i>et al.</i>	(OPAL Collab.)
DECAMP	91J	PL B266 218	D. Decamp <i>et al.</i>	(ALEPH Collab.)
ABACHI	88B	PL B212 533	S. Abachi <i>et al.</i>	(ANL, IND, MICH, PURD+)
ADLER	88D	PL B208 152	J. Adler <i>et al.</i>	(Mark III Collab.)
ALBRECHT	85F	PL 150B 235	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AHLEN	83	PRL 51 1147	S.P. Ahlen <i>et al.</i>	(ANL, IND, LBL+)
BAILEY	83	PL 132B 230	R. Bailey <i>et al.</i>	(AMST, BRIS, CERN, CRAC+)
COLES	82	PR D26 2190	M.W. Coles <i>et al.</i>	(LBL, SLAC)
YELTON	82	PRL 49 430	J.M. Yelton <i>et al.</i>	(SLAC, LBL, UCB+)
FITCH	81	PRL 46 761	V.L. Fitch <i>et al.</i>	(PRIN, SACL, TORI+)
AVERY	80	PRL 44 1309	P. Avery <i>et al.</i>	(ILL, FNAL, COLU)
BLIETSCHAU	79	PL B6B 108	J. Blietschau <i>et al.</i>	(AACH3, BONN, CERN+)
FELDMAN	77B	PRL 38 1313	G.J. Feldman <i>et al.</i>	(Mark I Collab.)
GOLDHABER	77	PL 69B 503	G. Goldhaber <i>et al.</i>	(Mark I Collab.)
PERUZZI	77	PRL 39 1301	I. Peruzzi <i>et al.</i>	(LGW Collab.)

## $D_0^*(2400)^0$

$J^P = 0^+$  assignment favored (ABE 04D).

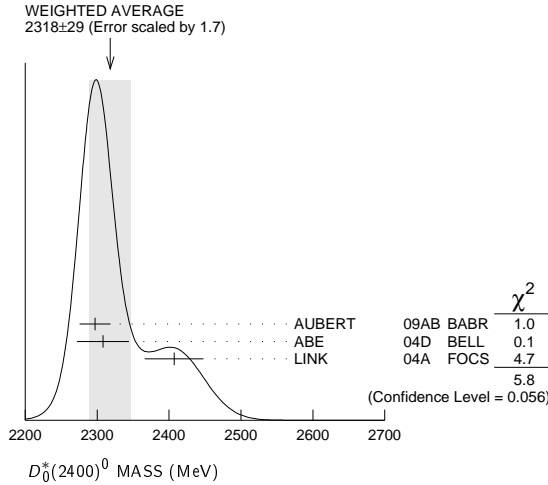
### $D_0^*(2400)^0$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2318±29 OUR AVERAGE</b> Error includes scale factor of 1.7. See the ideogram below.				
2297± 8±20	3.4k	AUBERT	09AB BABR	$B^- \rightarrow D^+ \pi^- \pi^-$
2308±17±32		ABE	04D BELL	$B^- \rightarrow D^+ \pi^- \pi^-$
2407±21±35	9.8k	LINK	04A FOCS	$\gamma A$



See key on page 601

## Meson Particle Listings

 $D_0^*(2400)^0, D_0^*(2400)^\pm, D_1(2420)^0$  $D_0^*(2400)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>267±40 OUR AVERAGE</b>				
273±12±48	3.4k	AUBERT	09AB BABR	$B^- \rightarrow D^+ \pi^- \pi^-$
276±21±63		ABE	04D BELL	$B^- \rightarrow D^+ \pi^- \pi^-$
240±55±59	9.8k	LINK	04A FOCS	$\gamma A$

 $D_0^*(2400)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^+ \pi^-$	seen

 $D_0^*(2400)^0$  REFERENCES

AUBERT	09AB PR D79 112004	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE	04D PR D69 112002	K. Abe <i>et al.</i>	(BELLE Collab.)
LINK	04A PL B586 11	J.M. Link <i>et al.</i>	(FOCUS Collab.)

 $D_0^*(2400)^\pm$ 

$$I(J^P) = \frac{1}{2}(0^+)$$

OMITTED FROM SUMMARY TABLE  
 $J, P$  need confirmation. $D_0^*(2400)^\pm$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2351±7 OUR AVERAGE</b>				
2360±15±30		1 AAIJ	15X LHCb	$B^0 \rightarrow \bar{D}^0 K^+ \pi^-$
2349±6±4		2 AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
2403±14±35	18.8k	LINK	04A FOCS	$\gamma A$
2354±7±11		3 AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> From the Dalitz plot analysis including various  $K^*$  and  $D^{**}$  mesons as well as broad structures in the  $K\pi$  S-wave and the  $D\pi$  S- and P-waves.  
<sup>2</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the Isobar formalism.  
<sup>3</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the K-matrix formalism.

 $D_0^*(2400)^\pm$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>230±17 OUR AVERAGE</b>				Error includes scale factor of 1.1.
255±26±51		1 AAIJ	15X LHCb	$B^0 \rightarrow \bar{D}^0 K^+ \pi^-$
217±13±13		2 AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
283±24±34	18.8k	LINK	04A FOCS	$\gamma A$
230±15±21		3 AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> From the Dalitz plot analysis including various  $K^*$  and  $D^{**}$  mesons as well as broad structures in the  $K\pi$  S-wave and the  $D\pi$  S- and P-waves.  
<sup>2</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the Isobar formalism.  
<sup>3</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the K-matrix formalism.

 $D_0^*(2400)^\pm$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^0 \pi^+$	seen

 $D_0^*(2400)^\pm$  REFERENCES

AAIJ	15X PR D92 012012	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15Y PR D92 032002	R. Aaij <i>et al.</i>	(LHCb Collab.)
LINK	04A PL B586 11	J.M. Link <i>et al.</i>	(FOCUS Collab.)

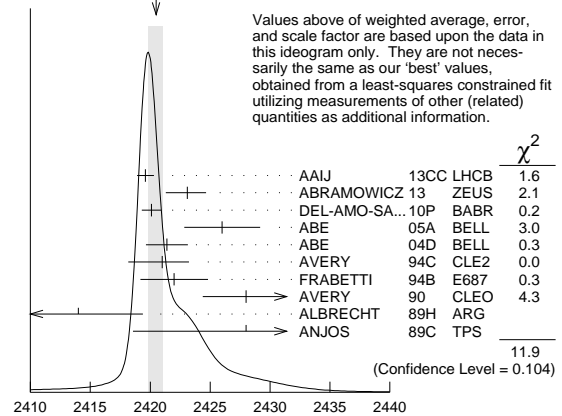
 $D_1(2420)^0$ 

$$I(J^P) = \frac{1}{2}(1^+)$$

$I$  needs confirmation.

 $D_1(2420)^0$  MASSThe fit includes  $D^\pm, D^0, D_s^\pm, D^{*+}, D^{*0}, D_s^{*+}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2420.8±0.5 OUR FIT</b>				Error includes scale factor of 1.3.
<b>2420.5±0.6 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
2419.6±0.1±0.7	210k	AAIJ	13cc LHCb	$p p \rightarrow D^{*+} \pi^- X$
2423.1±1.5±0.4	2.7k	1 ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)} \pi^- X$
2420.1±0.1±0.8	103k	DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^{*+} \pi^- X$
2426 ±3 ±1	151	ABE	05A BELL	$B^- \rightarrow D^0 \pi^+ \pi^- \pi^-$
2421.4±1.5±0.9		2 ABE	04D BELL	$B^- \rightarrow D^{*+} \pi^- \pi^-$
2421 ±1 ±2	286	AVERY	94C CLE2	$e^+ e^- \rightarrow D^{*+} \pi^- X$
2422 ±2 ±2	51	FRABETTI	94B E687	$\gamma Be \rightarrow D^{*+} \pi^- X$
2428 ±3 ±2	279	AVERY	90 CLEO	$e^+ e^- \rightarrow D^{*+} \pi^- X$
2414 ±2 ±5	171	ALBRECHT	89H ARG	$e^+ e^- \rightarrow D^{*+} \pi^- X$
2428 ±8 ±5	171	ANJOS	89C TPS	$\gamma N \rightarrow D^{*+} \pi^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2420.5±2.1±0.9	3110±340	3 CHEKANOV	09 ZEUS	$e^\pm p \rightarrow D^{*+} \pi^- X$
2421.7±0.7±0.6	7.5k	ABULENCIA	06A CDF	1900 $p\bar{p} \rightarrow D^{*+} \pi^- X$
2425 ±3 ±	235	4 ABREU	98M DLPH	$e^+ e^-$

<sup>1</sup> From the combined fit of the  $M(D^+ \pi^-)$  and  $M(D^{*+} \pi^-)$  distributions. and  $A_{D_2}$  fixed to the theoretical prediction of  $-1$ .<sup>2</sup> Fit includes the contribution from  $D_1^*(2430)^0$ .<sup>3</sup> Calculated using the mass difference  $m(D_1^0) - m(D^{*+})_{PDG}$  reported below and  $m(D^{*+})_{PDG} = 2010.27 \pm 0.17$  MeV. The 0.17 MeV uncertainty of the PDG mass value should be added to the experimental uncertainty of 0.9 MeV.<sup>4</sup> No systematic error given.WEIGHTED AVERAGE  
2420.5±0.6 (Error scaled by 1.3) $m_{D_1^0} - m_{D^{*+}}$ The fit includes  $D^\pm, D^0, D_s^\pm, D^{*+}, D^{*0}, D_s^{*+}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>410.6±0.5 OUR FIT</b>				Error includes scale factor of 1.3.
<b>411.5±0.8 OUR AVERAGE</b>				
410.2±2.1±0.9	3110±340	CHEKANOV	09 ZEUS	$e^\pm p \rightarrow D^{*+} \pi^- X$
411.7±0.7±0.4	7.5k	ABULENCIA	06A CDF	1900 $p\bar{p} \rightarrow D^{*+} \pi^- X$

# Meson Particle Listings

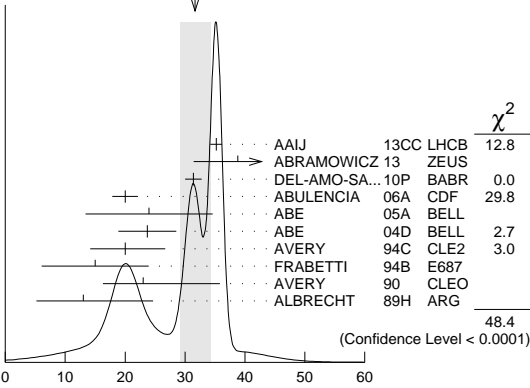
## $D_1(2420)^0, D_1(2420)^\pm$

### $D_1(2420)^0$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>31.7±2.5 OUR AVERAGE</b>		Error includes scale factor of 3.5. See the ideogram below.		
35.2±0.4±0.9	210k	AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^{*+}\pi^-X$
38.8±5.0±1.9	2.7k	<sup>1</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)}+\pi^-X$
31.4±0.5±1.3	103k	DEL-AMO-SA...10P	BABR	$e^+e^- \rightarrow D^{*+}\pi^-X$
20.0±1.7±1.3	7.5k	ABULENCIA	06A CDF	$1900\,p\bar{p} \rightarrow D^{*+}\pi^-X$
24±7±8	151	ABE	05A BELL	$B^- \rightarrow D^0\pi^+\pi^-\pi^-$
23.7±2.7±4.0		<sup>2</sup> ABE	04D BELL	$B^- \rightarrow D^{*+}\pi^-\pi^-$
20± <sup>6</sup> / <sub>5</sub> ±3	286	AVERY	94C CLE2	$e^+e^- \rightarrow D^{*+}\pi^-X$
15±8±4	51	FRABETTI	94B E687	$\gamma Be \rightarrow D^{*+}\pi^-X$
23± <sup>8</sup> / <sub>6</sub> ± <sup>10</sup> / <sub>3</sub>	279	AVERY	90 CLEO	$e^+e^- \rightarrow D^{*+}\pi^-X$
13±6± <sup>10</sup> / <sub>5</sub>	171	ALBRECHT	89H ARG	$e^+e^- \rightarrow D^{*+}\pi^-X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
53.2±7.2± <sup>3.3</sup> / <sub>4.9</sub>	3110±340	CHEKANOV	09 ZEUS	$e^\pm p \rightarrow D^{*+}\pi^-X$
58±14±10	171	ANJOS	89c TPS	$\gamma N \rightarrow D^{*+}\pi^-X$

<sup>1</sup> From the combined fit of the  $M(D^+\pi^-)$  and  $M(D^{*+}\pi^-)$  distributions. and  $A_{D_2}$  fixed to the theoretical prediction of  $-1$ .  
<sup>2</sup> Fit includes the contribution from  $D_1^*(2430)^0$ .

WEIGHTED AVERAGE  
31.7±2.5 (Error scaled by 3.5)



$D_1(2420)^0$  WIDTH (MeV)

### $D_1(2420)^0$ DECAY MODES

$\overline{D}_1(2420)^0$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^*(2010)^+\pi^-$	seen
$\Gamma_2$ $D^0\pi^+\pi^-$	seen
$\Gamma_3$ $D^0\rho^0$	
$\Gamma_4$ $D^0f_0(500)$	
$\Gamma_5$ $D_0^*(2400)^+\pi^-$	
$\Gamma_6$ $D^+\pi^-$	not seen
$\Gamma_7$ $D^{*0}\pi^+\pi^-$	not seen

### $D_1(2420)^0$ BRANCHING RATIOS

$\Gamma(D^*(2010)^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	ACKERSTAFF	97W OPAL	$e^+e^- \rightarrow D^{*+}\pi^-X$	
seen	AVERY	90 CLEO	$e^+e^- \rightarrow D^{*+}\pi^-X$	
seen	ALBRECHT	89H ARG	$e^+e^- \rightarrow D^*\pi^-X$	
seen	ANJOS	89c TPS	$\gamma N \rightarrow D^{*+}\pi^-X$	
$\Gamma(D^+\pi^-)/\Gamma(D^*(2010)^+\pi^-)$				$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.24	90	AVERY	90 CLEO	$e^+e^- \rightarrow D^+\pi^-X$

### $D_1(2420)^0$ POLARIZATION AMPLITUDE $A_{D_1}$

A polarization amplitude  $A_{D_1}$  is a parameter that depends on the initial polarization of the  $D_1$  and is sensitive to a possible S-wave contribution to its decay. For  $D_1$  decays the helicity angle,  $\theta_h$ , distribution varies like  $1 + A_{D_1}\cos^2\theta_h$ , where  $\theta_h$  is the angle in the  $D^*$  rest frame between the two pions emitted by the  $D_1 \rightarrow D^*\pi$  and the  $D^* \rightarrow D\pi$ .

Unpolarized  $D_1$  decaying purely via D-wave is predicted to give  $A_{D_1} = 3$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.73±0.25 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.		
7.8± <sup>6.7</sup> / <sub>2.7</sub> ± <sup>4.6</sup> / <sub>1.8</sub>	2.7k	<sup>1</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)}+\pi^-X$
5.72±0.25	103k	DEL-AMO-SA...10P	BABR	$e^+e^- \rightarrow D^{*+}\pi^-X$
5.9± <sup>3.0</sup> / <sub>1.7</sub> ± <sup>2.4</sup> / <sub>1.0</sub>		CHEKANOV	09 ZEUS	$e^\pm p \rightarrow D^{*+}\pi^-X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.30±0.48	210k	<sup>2</sup> AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^{*+}\pi^-X$
3.8±0.6±0.8		<sup>3</sup> AUBERT	09Y BABR	$B^+ \rightarrow D_1^0\ell^+\nu_\ell$
2.74± <sup>1.40</sup> / <sub>0.93</sub>		<sup>4</sup> AVERY	94C CLE2	$e^+e^- \rightarrow D^{*+}\pi^-X$

<sup>1</sup> From the combined fit of the  $M(D^+\pi^-)$  and  $M(D^{*+}\pi^-)$  distributions. and  $A_{D_2}$  fixed to the theoretical prediction of  $-1$ . A pure D-wave not excluded although some S-wave mixing possible.  
<sup>2</sup> Systematic uncertainty not estimated. Resonance parameters fixed.  
<sup>3</sup> Assuming  $\Gamma(\Upsilon(4S) \rightarrow B^+B^-) / \Gamma(\Upsilon(4S) \rightarrow B^0\overline{B}^0) = 1.065 \pm 0.026$  and equal partial widths and helicity angle distributions for charged and neutral  $D_1$  mesons.  
<sup>4</sup> Systematic uncertainties not estimated.

### $D_1(2420)^0$ REFERENCES

AAIJ	13CC JHEP 1309 145	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABRAMOWICZ 13	NP B866 229	H. Abramowicz <i>et al.</i>	(ZEUS Collab.)
DEL-AMO-SA...10P	PR D82 111101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
AUBERT	09Y PRL 103 051803	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHEKANOV	09Y EPJ C60 25	S. Chekanov <i>et al.</i>	(ZEUS Collab.)
ABULENCIA	06A PR D73 051104	A. Abulencia <i>et al.</i>	(CDF Collab.)
ABE	05A PRL 94 221805	K. Abe <i>et al.</i>	(BELLE Collab.)
ABE	04D PR D69 112002	K. Abe <i>et al.</i>	(BELLE Collab.)
ABREU	98M PL B426 231	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	97W ZPHY. C76 425	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
AVERY	94C PL B331 236	P. Avery <i>et al.</i>	(CLEO Collab.)
FRABETTI	94B PRL 72 324	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
AVERY	90 PR D41 774	P. Avery, D. Besson	(CLEO Collab.)
ALBRECHT	89H PL B232 398	H. Albrecht <i>et al.</i>	(ARGUS Collab.)JP
ANJOS	89C PRL 62 1717	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)

## $D_1(2420)^\pm$

$I(J^P) = \frac{1}{2}(?^?)$   
 $I$  needs confirmation.

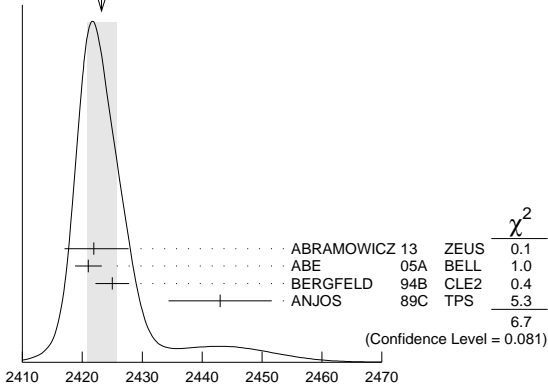
OMITTED FROM SUMMARY TABLE  
Seen in  $D^*(2007)^0\pi^+$ .  $J^P = 0^+\pi^0$  ruled out.

### $D_1(2420)^\pm$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2423.2±2.4 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.		
2421.9±4.7± <sup>3.4</sup> / <sub>1.2</sub>	759	<sup>1</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)0}\pi^+X$
2421±2±1	124	ABE	05A BELL	$\overline{B}^0 \rightarrow D^+\pi^+\pi^-\pi^-$
2425±2±2	146	BERGFELD	94B CLE2	$e^+e^- \rightarrow D^{*0}\pi^+X$
2443±7±5	190	ANJOS	89c TPS	$\gamma N \rightarrow D^0\pi^+X^0$

<sup>1</sup> From the fit of the  $M(D^0\pi^+)$  distribution. The widths of the  $D_1^+$  and  $D_2^{*+}$  are fixed to 25 MeV and 37 MeV, and  $A_{D_1}$  and  $A_{D_2}$  are fixed to the theoretical predictions of 3 and  $-1$ , respectively.

WEIGHTED AVERAGE  
2423.2±2.4 (Error scaled by 1.5)



$D_1(2420)^\pm$  MASS (MeV)

### $m_{D_1^*(2420)^\pm} - m_{D_1^*(2420)^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4±<sup>2</sup>/<sub>3</sub>±3</b>	BERGFELD	94B CLE2	$e^+e^- \rightarrow \text{hadrons}$

See key on page 601

Meson Particle Listings

$D_1(2420)^\pm$ ,  $D_1(2430)^0$ ,  $D_2^*(2460)^0$

$D_1(2420)^\pm$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>25 \pm 6</math> OUR AVERAGE</b>				
$21 \pm 5 \pm 8$	124	ABE	05A BELL	$\bar{B}^0 \rightarrow D^+ \pi^+ \pi^- \pi^-$
$26^{+8}_{-7} \pm 4$	146	BERGFELD	94B CLE2	$e^+ e^- \rightarrow D^{*0} \pi^+ X$
$41 \pm 19 \pm 8$	190	ANJOS	89C TPS	$\gamma N \rightarrow D^0 \pi^+ X^0$

$D_1(2420)^\pm$  DECAY MODES

$D_1^*(2420)^-$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^*(2007)^0 \pi^+$	seen
$\Gamma_2$ $D^+ \pi^+ \pi^-$	seen
$\Gamma_3$ $D^+ \rho^0$	
$\Gamma_4$ $D^+ f_0(500)$	
$\Gamma_5$ $D_0^*(2400)^0 \pi^+$	
$\Gamma_6$ $D^0 \pi^+$	not seen
$\Gamma_7$ $D^{*+} \pi^+ \pi^-$	not seen

$D_1(2420)^\pm$  BRANCHING RATIOS

$\Gamma(D^*(2007)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	ANJOS	89C TPS	$\gamma N \rightarrow D^0 \pi^+ X^0$

$\Gamma(D^0 \pi^+)/\Gamma(D^*(2007)^0 \pi^+)$	$\Gamma_6/\Gamma_1$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.18	90	BERGFELD	94B CLE2	$e^+ e^- \rightarrow \text{hadrons}$

$D_1(2420)^\pm$  POLARIZATION AMPLITUDE  $A_{D_1}$

A polarization amplitude  $A_{D_1}$  is a parameter that depends on the initial polarization of the  $D_1$  and is sensitive to a possible S-wave contribution to its decay. For  $D_1$  decays the helicity angle,  $\theta_h$ , distribution varies like  $1 + A_{D_1} \cos^2 \theta_h$ , where  $\theta_h$  is the angle in the  $D^*$  rest frame between the two pions emitted by the  $D_1 \rightarrow D^* \pi$  and the  $D^* \rightarrow D \pi$ .

Unpolarized  $D_1$  decaying purely via D-wave is predicted to give  $A_{D_1} = 3$ .

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.8 \pm 0.6 \pm 0.8$	<sup>2</sup> AUBERT	09Y BABR	$B^0 \rightarrow D_1^- \ell^+ \nu_\ell$

<sup>2</sup> Assuming  $\Gamma(\Upsilon(4S) \rightarrow B^+ B^-) / \Gamma(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = 1.065 \pm 0.026$  and equal partial widths and helicity angle distributions for charged and neutral  $D_1$  mesons.

$D_1(2420)^\pm$  REFERENCES

ABRAMOWICZ 13	NP B866 229	H. Abramowicz <i>et al.</i>	(ZEUS Collab.)
AUBERT 09Y	PRL 103 051803	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE 05A	PRL 94 221805	K. Abe <i>et al.</i>	(BELLE Collab.)
BERGFELD 94B	PL B340 194	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
ANJOS 89C	PRL 62 1717	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)

$D_1(2430)^0$

$I(J^P) = \frac{1}{2}(1^+)$

OMITTED FROM SUMMARY TABLE  
 $J = 1^+$  assignment favored (ABE 04D).

$D_1(2430)^0$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>2427 \pm 26 \pm 25</math></b>	ABE	04D BELL	$B^- \rightarrow D^{*+} \pi^- \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2477 \pm 28$	<sup>1</sup> AUBERT	06L BABR	$\bar{B}^0 \rightarrow D^{*+} \omega \pi^-$

<sup>1</sup> Systematic errors not estimated.

$D_1(2430)^0$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>384^{+107}_{-75} \pm 74</math></b>	ABE	04D BELL	$B^- \rightarrow D^{*+} \pi^- \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$266 \pm 97$	<sup>2</sup> AUBERT	06L BABR	$\bar{B}^0 \rightarrow D^{*+} \omega \pi^-$

<sup>2</sup> Systematic errors not estimated.

$D_1(2430)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^*(2010)^+ \pi^-$	seen

$D_1(2430)^0$  REFERENCES

AUBERT 06L	PR D74 012001	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE 04D	PR D69 112002	K. Abe <i>et al.</i>	(BELLE Collab.)

$D_2^*(2460)^0$

$I(J^P) = \frac{1}{2}(2^+)$

$J^P = 2^+$  assignment strongly favored (ALBRECHT 89B, ALBRECHT 89H), natural parity confirmed by the helicity analysis (DEL-AMO-SANCHEZ 10P). AAIJ 13CC confirms  $J^P = 2^+$  and natural parity.

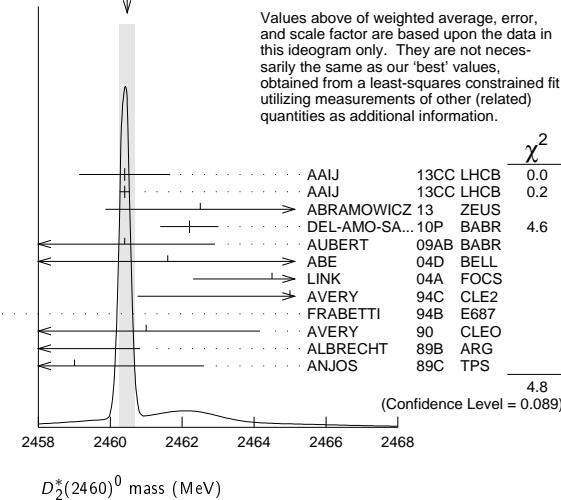
$D_2^*(2460)^0$  MASS

The fit includes  $D^\pm$ ,  $D^0$ ,  $D_s^\pm$ ,  $D^{*0}$ ,  $D_s^{*0}$ ,  $D_1(2420)^0$ ,  $D_2^*(2460)^0$ , and  $D_{S1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2460.57 \pm 0.15</math> OUR FIT</b>				Error includes scale factor of 1.1.
<b><math>2460.47 \pm 0.21</math> OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
$2460.4 \pm 0.4 \pm 1.2$	82k	AAIJ	13CC LHCB	$p p \rightarrow D^{*+} \pi^- X$
$2460.4 \pm 0.1 \pm 0.1$	675k	AAIJ	13CC LHCB	$p p \rightarrow D^+ \pi^- X$
$2462.5 \pm 2.4^{+1.3}_{-1.1}$	2.3k	<sup>1</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)+} \pi^- X$
$2462.2 \pm 0.1 \pm 0.8$	243k	DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^+ \pi^- X$
$2460.4 \pm 1.2 \pm 2.2$	3.4k	AUBERT 09AB	BABR	$B^- \rightarrow D^+ \pi^- \pi^-$
$2461.6 \pm 2.1 \pm 3.3$		<sup>2</sup> ABE 04D	BELL	$B^- \rightarrow D^+ \pi^- \pi^-$
$2464.5 \pm 1.1 \pm 1.9$	5.8k	<sup>2</sup> LINK 04A	FOCS	$\gamma A$
$2465 \pm 3 \pm 3$	486	AVERY 94C	CLE2	$e^+ e^- \rightarrow D^+ \pi^- X$
$2453 \pm 3 \pm 2$	128	FRABETTI 94B	E687	$\gamma \text{Be} \rightarrow D^+ \pi^- X$
$2461 \pm 3 \pm 1$	440	AVERY 90	CLEO	$e^+ e^- \rightarrow D^{*+} \pi^- X$
$2455 \pm 3 \pm 5$	337	ALBRECHT 89B	ARG	$e^+ e^- \rightarrow D^+ \pi^- X$
$2459 \pm 3 \pm 2$	153	ANJOS 89C	TPS	$\gamma N \rightarrow D^+ \pi^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2469.1 \pm 3.7^{+1.2}_{-1.3}$	1.5k	<sup>3</sup> CHEKANOV 09	ZEUS	$e^\pm p \rightarrow D^{(*)+} \pi^- X$
$2463.3 \pm 0.6 \pm 0.8$	20k	ABULENCIA 06A	CDF	$1900 p \bar{p} \rightarrow D^+ \pi^- X$
$2461 \pm 6$	126	<sup>4</sup> ABREU 98M	DLPH	$e^+ e^-$
$2466 \pm 7$	1	ASRATYAN 95	BEBC	$53,40 \nu(\bar{\nu}) \rightarrow p X, d X$

- <sup>1</sup> From the combined fit of the  $M(D^+ \pi^-)$  and  $M(D^{*+} \pi^-)$  distributions. and  $A_{D_2}$  fixed to the theoretical prediction of  $-1$ .
- <sup>2</sup> Fit includes the contribution from  $D_0^*(2400)^0$ .
- <sup>3</sup> Calculated using the mass difference  $m(D_2^{*0}) - m(D^{*+})_{PDG}$  reported below and  $m(D^{*+})_{PDG} = 2010.27 \pm 0.17$  MeV. The 0.17 MeV uncertainty of the PDG mass value should be added to the experimental uncertainty of  $^{+1.2}_{-1.3}$  MeV.
- <sup>4</sup> No systematic error given.

WEIGHTED AVERAGE  
 $2460.47 \pm 0.21$  (Error scaled by 1.6)



Meson Particle Listings

$D_2^*(2460)^0$

$m_{D_2^{*0}} - m_{D^+}$

The fit includes  $D^\pm, D^0, D_S^\pm, D^{*\pm}, D^{*0}, D_S^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{S1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>590.98±0.18 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>593.9 ± 0.6 ± 0.5</b>	20k	ABULENCIA	06A CDF	1900 $p\bar{p} \rightarrow D^+\pi^-X$

$m_{D_2^{*0}} - m_{D^{*+}}$

The fit includes  $D^\pm, D^0, D_S^\pm, D^{*\pm}, D^{*0}, D_S^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{S1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>450.31±0.16 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>458.8 ± 3.7 <sup>+1.2</sup><sub>-1.3</sub></b>	1560±230	CHEKANOV	09 ZEUS	$e^\pm p \rightarrow D^{(*)+}\pi^-X$

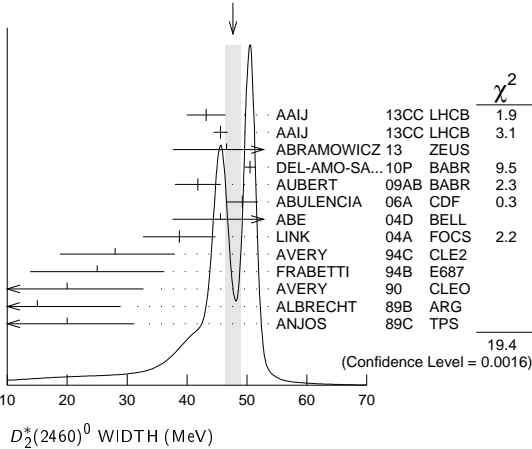
$D_2^*(2460)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>47.7± 1.3 OUR AVERAGE</b>	Error includes scale factor of 2.0. See the ideogram below.			
43.2± 1.2± 3.0	82k	AAIJ	13CC LHCB	$pp \rightarrow D^{*+}\pi^-X$
45.6± 0.4± 1.1	675k	AAIJ	13CC LHCB	$pp \rightarrow D^+\pi^-X$
46.6± 8.1 <sup>+</sup> <sub>-5.9</sub> 3.8	2.3k	<sup>5</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)+}\pi^-X$
50.5± 0.6± 0.7	243k	DEL-AMO-SA...10P	BABR	$e^+e^- \rightarrow D^+\pi^-X$
41.8± 2.5± 2.9	3.4k	AUBERT	09AB BABR	$B^- \rightarrow D^+\pi^-\pi^-$
49.2± 2.3± 1.3	20k	ABULENCIA	06A CDF	1900 $p\bar{p} \rightarrow D^+\pi^-X$
45.6± 4.4± 6.7		<sup>6</sup> ABE	04D BELL	$B^- \rightarrow D^+\pi^-\pi^-$
38.7± 5.3± 2.9	5.8k	<sup>6</sup> LINK	04A FOCUS	$\gamma A$
28 <sup>+</sup> <sub>-7</sub> ± 6	486	AVERY	94C CLE2	$e^+e^- \rightarrow D^+\pi^-X$
25 ±10 ± 5	128	FRABETTI	94B E687	$\gamma\text{Be} \rightarrow D^+\pi^-X$
20 <sup>+</sup> <sub>-12</sub> <sup>9</sup> <sub>-10</sub>	440	AVERY	90 CLEO	$e^+e^- \rightarrow D^{*+}\pi^-X$
15 <sup>+13</sup> <sub>-10</sub> <sup>5</sup> <sub>-10</sub>	337	ALBRECHT	89B ARG	$e^+e^- \rightarrow D^+\pi^-X$
20 ±10 ± 5	153	ANJOS	89C TPS	$\gamma N \rightarrow D^+\pi^-X$

<sup>5</sup> From the combined fit of the  $M(D^+\pi^-)$  and  $M(D^{*+}\pi^-)$  distributions. and  $A_{D_2}$  fixed to the theoretical prediction of  $-1$ .

<sup>6</sup> Fit includes the contribution from  $D_0^*(2400)^0$ .

WEIGHTED AVERAGE  
47.7±1.3 (Error scaled by 2.0)



$D_2^*(2460)^0$  DECAY MODES

$\bar{D}_2^*(2460)^0$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^+\pi^-$	seen
$\Gamma_2$ $D^{*}(2010)^+\pi^-$	seen
$\Gamma_3$ $D^0\pi^+\pi^-$	not seen
$\Gamma_4$ $D^{*0}\pi^+\pi^-$	not seen

$D_2^*(2460)^0$  BRANCHING RATIOS

$\Gamma(D^+\pi^-)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
seen	3.4k		AUBERT	09AB BABR	$B^- \rightarrow D^+\pi^-\pi^-$	
seen	337		ALBRECHT	89B ARG	$e^+e^- \rightarrow D^+\pi^-X$	
seen			ANJOS	89C TPS	$\gamma N \rightarrow D^+\pi^-X$	

$\Gamma(D^{*}(2010)^+\pi^-)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
seen			ACKERSTAFF	97W OPAL	$e^+e^- \rightarrow D^{*+}\pi^-X$	
seen			AVERY	90 CLEO	$e^+e^- \rightarrow D^{*+}\pi^-X$	
seen			ALBRECHT	89H ARG	$e^+e^- \rightarrow D^{*+}\pi^-X$	

$\Gamma(D^+\pi^-)/\Gamma(D^{*}(2010)^+\pi^-)$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma_2$
<b>1.54±0.15 OUR AVERAGE</b>						
1.4 ± 0.3 ± 0.3	2.3k	<sup>7</sup> ABRAMOWICZ13	ZEUS		$e^\pm p \rightarrow D^{(*)+}\pi^-X$	
1.47±0.03±0.16	379k	DEL-AMO-SA...10P	BABR		$e^+e^- \rightarrow D^{(*)+}\pi^-X$	
2.8 ± 0.8 <sup>+</sup> <sub>-0.6</sub>	1560±230	CHEKANOV	09 ZEUS		$e^\pm p \rightarrow D^{(*)+}\pi^-X$	
2.2 ± 0.7 ± 0.6		AVERY	94C CLE2		$e^+e^- \rightarrow D^{*+}\pi^-X$	
2.3 ± 0.8		AVERY	90 CLEO		$e^+e^- \rightarrow D^{*+}\pi^-X$	
3.0 ± 1.1 ± 1.5		ALBRECHT	89H ARG		$e^+e^- \rightarrow D^{*+}\pi^-X$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1.9 ± 0.5		ABE	04D BELL		$B^- \rightarrow D^{(*)+}\pi^-X$	

<sup>7</sup> From the combined fit of the  $M(D^+\pi^-)$  and  $M(D^{*+}\pi^-)$  distributions. and  $A_{D_2}$  fixed to the theoretical prediction of  $-1$ .

$\Gamma(D^+\pi^-)/[\Gamma(D^+\pi^-) + \Gamma(D^{*}(2010)^+\pi^-)]$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/(\Gamma_1+\Gamma_2)$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.62±0.03±0.02	8414	<sup>8</sup> AUBERT	09Y BABR		$B^+ \rightarrow D_2^{*0}\ell^+\nu_\ell$	

<sup>8</sup> Assuming  $\Gamma(\Upsilon(4S) \rightarrow B^+B^-) / \Gamma(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = 1.065 \pm 0.026$  and equal partial widths for charged and neutral  $D_2^*$  mesons.

$D_2^*(2460)^0$  POLARIZATION AMPLITUDE  $A_{D_2}$

A polarization amplitude  $A_{D_2}$  is a parameter that depends on the initial polarization of the  $D_2$ . For  $\bar{D}_2$  decays the helicity angle,  $\theta_H$ , distribution varies like  $1 + A_{D_2} \cos^2(\theta_H)$ , where  $\theta_H$  is the angle in the  $D^*$  rest frame between the two pions emitted by the  $D_2 \rightarrow D^*\pi$  and  $D^* \rightarrow D\pi$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-1.16 \pm 0.35$	2.3k	<sup>9</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)+}\pi^-X$
consistent with $-1$	243k	DEL-AMO-SA...10P	BABR	$e^+e^- \rightarrow D^+\pi^-X$
$-0.74 \pm 0.49$ $-0.38$		<sup>10</sup> AVERY	94C CLE2	$e^+e^- \rightarrow D^{*+}\pi^-X$

<sup>9</sup> From the combined fit of the  $M(D^+\pi^-)$  and  $M(D^{*+}\pi^-)$  distributions.

<sup>10</sup> Systematic uncertainties not estimated.

$D_2^*(2460)^0$  REFERENCES

AAIJ	13CC JHEP 1309 145	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABRAMOWICZ 13	NP B866 229	H. Abramowicz <i>et al.</i>	(ZEUS Collab.)
DEL-AMO-SA...10P	PR D82 111101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
AUBERT	09AB PR D79 112004	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09Y PRL 103 051803	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHEKANOV	09 EPJ C60 25	S. Chekanov <i>et al.</i>	(ZEUS Collab.)
ABULENCIA	06A PR D73 051104	A. Abulencia <i>et al.</i>	(CDF Collab.)
ABE	04D PR D69 112002	K. Abe <i>et al.</i>	(BELLE Collab.)
LINK	04A PL B586 11	J.M. Link <i>et al.</i>	(FOCUS Collab.)
ABREU	98M PL B426 231	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	97W ZPHY C76 425	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ASRATYAN	95 ZPHY C68 43	A.E. Asratyan <i>et al.</i>	(BIRM, BELG, CERN+)
AVERY	94C PL B331 236	P. Avery <i>et al.</i>	(CLEO Collab.)
FRABETTI	94B PR 72 324	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
AVERY	90 PR D41 774	P. Avery, D. Besson	(CLEO Collab.)
ALBRECHT	89B PL B221 422	H. Albrecht <i>et al.</i>	(ARGUS Collab.)JP
ALBRECHT	89H PL B232 398	H. Albrecht <i>et al.</i>	(ARGUS Collab.)JP
ANJOS	89C PRL 62 1717	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)

See key on page 601

## Meson Particle Listings

$$D_2^*(2460)^\pm, D(2550)^0$$

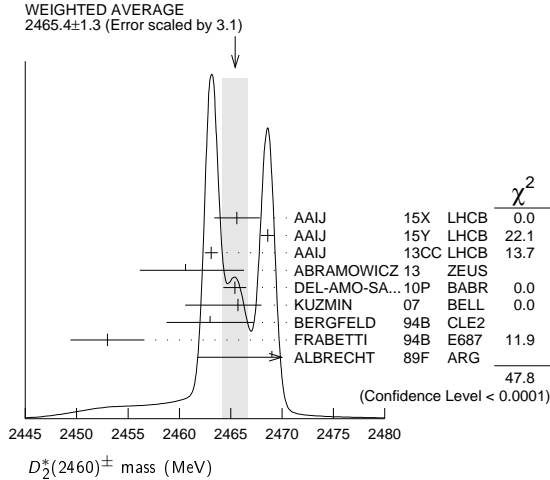
$$D_2^*(2460)^\pm$$

$$I(J^P) = \frac{1}{2}(2^+)$$

 $J^P = 2^+$  assignment strongly favored(ALBRECHT 89B).
 $D_2^*(2460)^\pm$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2465.4±1.3 OUR AVERAGE</b>		Error includes scale factor of 3.1. See the ideogram below.		
2465.6±1.8±1.3		<sup>1</sup> AAIJ	15X LHCb	$B^0 \rightarrow \bar{D}^0 K^+ \pi^-$
2468.6±0.6±0.3		<sup>2</sup> AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
2463.1±0.2±0.6	342k	AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^0 \pi^+ X$
2460.6±4.4+3.6-0.8	1371	<sup>3</sup> ABRAMOWICZ13	ZEUS	$e^\pm p \rightarrow D^{(*)0} \pi^+ X$
2465.4±0.2±1.1	111k	<sup>4</sup> DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^0 \pi^+ X$
2465.7±1.8+1.4-4.8	2909	KUZMIN	07 BELL	$e^+ e^- \rightarrow \text{hadrons}$
2463 ±3 ±3	310	BERGFELD	94B CLE2	$e^+ e^- \rightarrow D^0 \pi^+ X$
2453 ±3 ±2	185	FRABETTI	94B E687	$\gamma \text{Be} \rightarrow D^0 \pi^+ X$
2469 ±4 ±6		ALBRECHT	89F ARG	$e^+ e^- \rightarrow D^0 \pi^+ X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2468.1±0.6±0.5		<sup>5</sup> AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
2467.6±1.5±0.8	3.5k	<sup>6</sup> LINK	04A FOCUS	$\gamma A$

- <sup>1</sup> From the Dalitz plot analysis including various  $K^*$  and  $D^{**}$  mesons as well as broad structures in the  $K\pi$  S-wave and the  $D\pi$  S- and P-waves.  
<sup>2</sup> Modeling the  $\pi^+\pi^-$  S-wave with the isobar formalism.  
<sup>3</sup> From the fit of the  $M(D^0\pi^+)$  distribution. The widths of the  $D_1^+$  and  $D_2^{*+}$  are fixed to 25 MeV and 37 MeV, and  $A_{D_1}$  and  $A_{D_2}$  are fixed to the theoretical predictions of 3 and -1, respectively.  
<sup>4</sup> At a fixed width of 50.5 MeV.  
<sup>5</sup> Modeling the  $\pi^+\pi^-$  S-wave with the K-matrix formalism.  
<sup>6</sup> Fit includes the contribution from  $D_0^*(2400)^\pm$ . Not independent of the corresponding mass difference measurement,  $(m_{D_2^*(2460)^\pm} - m_{D_2^*(2460)^0})$ .



$$m_{D_2^*(2460)^\pm} - m_{D_2^*(2460)^0}$$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2.4±1.7 OUR AVERAGE</b>			
3.1±1.9±0.9	LINK	04A FOCUS	$\gamma A$
-2 ±4 ±4	BERGFELD	94B CLE2	$e^+ e^- \rightarrow \text{hadrons}$
0 ±4	FRABETTI	94B E687	$\gamma \text{Be} \rightarrow D\pi X$
14 ±5 ±8	ALBRECHT	89F ARG	$e^+ e^- \rightarrow D^0 \pi^+ X$

 $D_2^*(2460)^\pm$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>46.7±1.2 OUR AVERAGE</b>				
46.0±3.4±3.2		<sup>1</sup> AAIJ	15X LHCb	$B^0 \rightarrow \bar{D}^0 K^+ \pi^-$
47.3±1.5±0.7		<sup>2</sup> AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
48.6±1.3±1.9	342k	AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^0 \pi^+ X$
49.7±3.8±6.4	2909	KUZMIN	07 BELL	$e^+ e^- \rightarrow \text{hadrons}$
34.1±6.5±4.2	3.5k	<sup>3</sup> LINK	04A FOCUS	$\gamma A$
27 ±11 ±8	310	BERGFELD	94B CLE2	$e^+ e^- \rightarrow D^0 \pi^+ X$
23 ±9 ±5	185	FRABETTI	94B E687	$\gamma \text{Be} \rightarrow D^0 \pi^+ X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
46.0±1.4±1.8		<sup>4</sup> AAIJ	15Y LHCb	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$

- <sup>1</sup> From the Dalitz plot analysis including various  $K^*$  and  $D^{**}$  mesons as well as broad structures in the  $K\pi$  S-wave and the  $D\pi$  S- and P-waves.  
<sup>2</sup> Modeling the  $\pi^+\pi^-$  S-wave with the isobar formalism.  
<sup>3</sup> Fit includes the contribution from  $D_0^*(2400)^\pm$ .  
<sup>4</sup> Modeling the  $\pi^+\pi^-$  S-wave with the K-matrix formalism.

 $D_2^*(2460)^\pm$  DECAY MODES
 $D_2^*(2460)^-$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^0 \pi^+$	seen
$\Gamma_2$ $D^{*0} \pi^+$	seen
$\Gamma_3$ $D^+ \pi^+ \pi^-$	not seen
$\Gamma_4$ $D^{*+} \pi^+ \pi^-$	not seen

 $D_2^*(2460)^\pm$  BRANCHING RATIOS

$\Gamma(D^0\pi^+)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	ALBRECHT	89F ARG	$e^+e^- \rightarrow D^0\pi^+X$
$\Gamma(D^0\pi^+)/\Gamma(D^{*0}\pi^+)$	$\Gamma_1/\Gamma_2$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.2±0.4 OUR AVERAGE</b>			
$1.1 \pm 0.4^{+0.3}_{-0.2}$	1371	<sup>1</sup> ABRAMOWICZ13	ZEUS $e^\pm p \rightarrow D^{(*)0}\pi^+X$
$1.9 \pm 1.1 \pm 0.3$		BERGFELD 94B CLE2	$e^+e^- \rightarrow \text{hadrons}$

<sup>1</sup> From the fit of the  $M(D^0\pi^+)$  distribution. The widths of the  $D_1^+$  and  $D_2^{*+}$  are fixed to 25 MeV and 37 MeV, and  $A_{D_1}$  and  $A_{D_2}$  are fixed to the theoretical predictions of 3 and -1, respectively.

$\Gamma(D^0\pi^+)/[\Gamma(D^0\pi^+)+\Gamma(D^{*0}\pi^+)]$					$\Gamma_1/(\Gamma_1+\Gamma_2)$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.62±0.03±0.02	3361	<sup>1</sup> AUBERT	09Y BABR	$\bar{B}^0 \rightarrow D_2^{*+} \ell^- \nu_\ell$	
<sup>1</sup> Assuming $\Gamma(\Upsilon(4S) \rightarrow B^+ B^-) / \Gamma(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = 1.065 \pm 0.026$ and equal partial widths for charged and neutral $D_2^*$ mesons.					

 $D_2^*(2460)^\pm$  REFERENCES

AAIJ	15X	PR D92 012012	R. Aaij et al.	(LHCb Collab.)
AAIJ	15Y	PR D92 032002	R. Aaij et al.	(LHCb Collab.)
AAIJ	13CC	JHEP 1309 145	R. Aaij et al.	(LHCb Collab.)
ABRAMOWICZ 13	NP	B866 229	H. Abramowicz et al.	(ZEUS Collab.)
DEL-AMO-SA...10P	PR	D82 111101	P. del Amo Sanchez et al.	(BABAR Collab.)
AUBERT	09Y	PRL 103 051803	B. Aubert et al.	(BABAR Collab.)
KUZMIN	07	PR D76 012006	A. Kuzmin et al.	(BELLE Collab.)
LINK	04A	PL B586 11	J.M. Link et al.	(FOCUS Collab.)
BERGFELD	94B	PL B340 194	T. Bergfeld et al.	(CLEO Collab.)
FRABETTI	94B	PRL 72 324	P.L. Frabetti et al.	(FNAL E687 Collab.)
ALBRECHT	89B	PL B221 422	H. Albrecht et al.	(ARGUS Collab.)
ALBRECHT	89F	PL B231 208	H. Albrecht et al.	(ARGUS Collab.)

$$D(2550)^0$$

$$I(J^P) = \frac{1}{2}(?^?)$$

OMITTED FROM SUMMARY TABLE

Unnatural parity according to the helicity analysis of DEL-AMO-SANCHEZ 10P and AAIJ 13CC. DEL-AMO-SANCHEZ 10P suggests  $J^P = 0^-$ .

 $D(2550)^0$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2564 ±20 OUR AVERAGE</b>		Error includes scale factor of 3.9.		
2579.5±3.4±5.5	60k	AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^{*+} \pi^- X$
2539.4±4.5±6.8	34k	DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^{*+} \pi^- X$

 $D(2550)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>135 ±17 OUR AVERAGE</b>				
177.5±17.8±46.0	60k	AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^{*+} \pi^- X$
130 ±12 ±13	34k	DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^{*+} \pi^- X$

 $D(2550)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^{*+} \pi^-$	seen

 $D(2550)^0$  POLARIZATION AMPLITUDE  $A_{D_J}$ 

A polarization amplitude  $A_{D_J}$  is a parameter that depends on the initial polarization of the  $D_J$ . For  $D_J$  decays the helicity angle,  $\theta_H$ , distribution varies like  $1 + A_{D_J} \cos^2(\theta_H)$ , where  $\theta_H$  is the angle in the  $D_J$  rest frame between the two pions emitted in the  $D_J \rightarrow D^* \pi$  and  $D^* \rightarrow D\pi$  decays.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
4.2±1.3	60k	<sup>1</sup> AAIJ	13CC LHCb	$p\bar{p} \rightarrow D^{*+} \pi^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				

Meson Particle Listings

$D(2550)^0, D_J^*(2600), D^*(2640)^\pm, D(2740)^0$

<sup>1</sup> Systematic uncertainty not estimated.

$D(2550)^0$  REFERENCES

AAIJ	13CC	JHEP 1309 145	R. Aaij <i>et al.</i>	(LHCb Collab.)
DEL-AMO-SA...10P	PR	D82 111101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)

$D_J^*(2600)$

was  $D(2600)$ ,

$I(J^P) = \frac{1}{2}(?^?)$

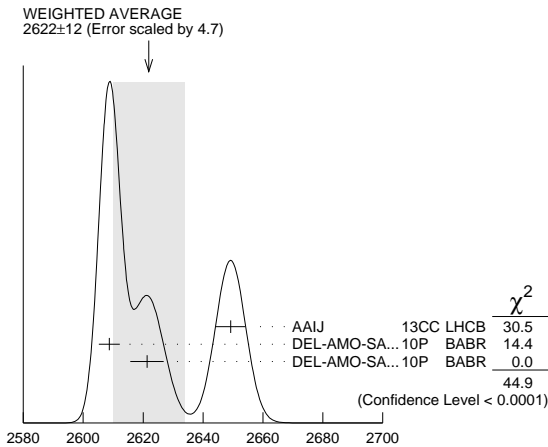
OMITTED FROM SUMMARY TABLE

$J^P$  consistent with natural parity (DEL-AMO-SANCHEZ 10P, AAIJ 13CC).

$D_J^*(2600)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2622 ±12</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 4.7. See the ideogram below.			
2649.2±3.5±3.5	51k	AAIJ	13cc	LHCB	$pp \rightarrow D^{*+} \pi^- X$
2608.7±2.4±2.5	26k	DEL-AMO-SA...10P	BABR	0	$e^+ e^- \rightarrow D^+ \pi^- X$
2621.3±3.7±4.2	13k	<sup>1</sup> DEL-AMO-SA...10P	BABR	+	$e^+ e^- \rightarrow D^0 \pi^+ X$

<sup>1</sup> At a fixed width of 93 MeV.



$D_J^*(2600)$  MASS (MeV)

$D_J^*(2600)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>104 ±20</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.6.		
140.2±17.1±18.6	51k	AAIJ	13cc	LHCB $pp \rightarrow D^{*+} \pi^- X$
93 ± 6 ±13	26k	DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^+ \pi^- X$

$D_J^*(2600)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D \pi$	seen
$\Gamma_2$ $D^+ \pi^-$	seen
$\Gamma_3$ $D^0 \pi^\pm$	seen
$\Gamma_4$ $D^* \pi$	seen
$\Gamma_5$ $D^{*+} \pi^-$	seen

$D_J^*(2600)$  BRANCHING RATIOS

$\Gamma(D^+ \pi^-)/\Gamma(D^{*+} \pi^-)$	$\Gamma_2/\Gamma_5$
<b>0.32±0.02±0.09</b>	
76k	
DEL-AMO-SA...10P	BABR
$e^+ e^- \rightarrow D(*)^+ \pi^- X$	

$D_J^*(2600)$  REFERENCES

AAIJ	13CC	JHEP 1309 145	R. Aaij <i>et al.</i>	(LHCb Collab.)
DEL-AMO-SA...10P	PR	D82 111101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)

$D^*(2640)^\pm$

$I(J^P) = \frac{1}{2}(?^?)$

OMITTED FROM SUMMARY TABLE

Seen in Z decays by ABREU 98M. Not seen by ABBIENDI 01N and CHEKANOV 09. Needs confirmation.

$D^*(2640)^\pm$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2637±2±6</b>	66 ± 14	ABREU	98M	DLPH $e^+ e^- \rightarrow D^{*+} \pi^+ \pi^- X$

$D^*(2640)^\pm$  WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;15</b>	95	ABREU	98M	DLPH $e^+ e^- \rightarrow D^{*+} \pi^+ \pi^- X$

$D^*(2640)^+$  DECAY MODES

$D^*(2640)^-$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^*(2010)^+ \pi^+ \pi^-$	seen

$D^*(2640)^\pm$  REFERENCES

CHEKANOV	09	EPJ C60 25	S. Chekanov <i>et al.</i>	(ZEUS Collab.)
ABBIENDI	01N	EPJ C20 445	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	98M	PL B426 231	P. Abreu <i>et al.</i>	(DELPHI Collab.)

$D(2740)^0$

$I(J^P) = \frac{1}{2}(?^?)$

OMITTED FROM SUMMARY TABLE

$J^P$  consistent with unnatural parity (AAIJ 13CC).

$D(2740)^0$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2737.0±3.5±11.2</b>	7.7k	AAIJ	13cc	LHCB $pp \rightarrow D^{*+} \pi^- X$

$D(2740)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>73.2±13.4±25.0</b>	7.7k	AAIJ	13cc	LHCB $pp \rightarrow D^{*+} \pi^- X$

$D(2740)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^{*+} \pi^-$	seen

$D(2740)^0$  POLARIZATION AMPLITUDE  $A_{D_J}$

A polarization amplitude  $A_{D_J}$  is a parameter that depends on the initial polarization of the  $D_J$ . For  $D_J$  decays the helicity angle,  $\theta_H$ , distribution varies like  $1 + A_{D_J} \cos^2(\theta_H)$ , where  $\theta_H$  is the angle in the  $D_J$  rest frame between the two pions emitted in the  $D_J \rightarrow D^* \pi$  and  $D^* \rightarrow D \pi$  decays.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
3.1±2.2	7.7k	<sup>1</sup> AAIJ	13cc	LHCB $pp \rightarrow D^{*+} \pi^- X$

<sup>1</sup> Systematic uncertainty not estimated.

$D(2740)^0$  REFERENCES

AAIJ	13CC	JHEP 1309 145	R. Aaij <i>et al.</i>	(LHCb Collab.)
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See key on page 601

## Meson Particle Listings

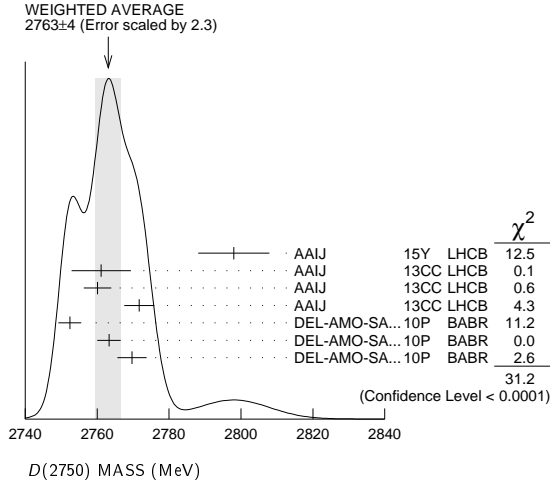
 $D(2750)$ ,  $D(3000)^0$  $D(2750)$ 

$$I(J^P) = \frac{1}{2}(3^-)$$

OMITTED FROM SUMMARY TABLE

 $J^P$  determined by AAIJ 15Y from the Dalitz plot analysis of  $B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$  decays.  $J^P$  consistent with natural parity (AAIJ 13CC). $D(2750)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2763 ± 4</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 2.3. See the ideogram below.			
2798 ± 7 ± 7		<sup>1</sup> AAIJ	15Y	LHCB	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
2761.1 ± 5.1 ± 6.5	14k	AAIJ	13CC	LHCB 0	$pp \rightarrow D^{*+} \pi^- X$
2760.1 ± 1.1 ± 3.7	56k	AAIJ	13CC	LHCB 0	$pp \rightarrow D^+ \pi^- X$
2771.7 ± 1.7 ± 3.8	20k	AAIJ	13CC	LHCB +	$pp \rightarrow D^0 \pi^+ X$
2752.4 ± 1.7 ± 2.7	23.5k	<sup>2</sup> DEL-AMO-SA...10P	BABR	0	$e^+ e^- \rightarrow D^{*+} \pi^- X$
2763.3 ± 2.3 ± 2.3	11.3k	<sup>2</sup> DEL-AMO-SA...10P	BABR	0	$e^+ e^- \rightarrow D^+ \pi^- X$
2769.7 ± 3.8 ± 1.5	5.7k	<sup>2,3</sup> DEL-AMO-SA...10P	BABR	+	$e^+ e^- \rightarrow D^0 \pi^+ X$
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
2802 ± 11 ± 10		<sup>4</sup> AAIJ	15Y	LHCB	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$

<sup>1</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the Isobar formalism.<sup>2</sup> The states observed in the  $D^* \pi$  and  $D \pi$  final states are not necessarily the same.<sup>3</sup> At a fixed width of 60.9 MeV.<sup>4</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the K-matrix formalism. $D(2750)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>65 ± 5</b>	<b>OUR AVERAGE</b>				
105 ± 18 ± 24		<sup>5</sup> AAIJ	15Y	LHCB	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$
74.4 ± 3.4 ± 37.0	14k	AAIJ	13CC	LHCB 0	$pp \rightarrow D^{*+} \pi^- X$
74.4 ± 3.4 ± 19.1	56k	AAIJ	13CC	LHCB 0	$pp \rightarrow D^+ \pi^- X$
66.7 ± 6.6 ± 10.5	20k	AAIJ	13CC	LHCB +	$pp \rightarrow D^0 \pi^+ X$
71 ± 6 ± 11	23.5k	<sup>6</sup> DEL-AMO-SA...10P	BABR		$e^+ e^- \rightarrow D^{*+} \pi^- X$
60.9 ± 5.1 ± 3.6	11.3k	<sup>6</sup> DEL-AMO-SA...10P	BABR		$e^+ e^- \rightarrow D^+ \pi^- X$
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
154 ± 27 ± 16		<sup>7</sup> AAIJ	15Y	LHCB	$B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-$

<sup>5</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the Isobar formalism.<sup>6</sup> The states observed in the  $D^* \pi$  and  $D \pi$  final states are not necessarily the same.<sup>7</sup> Modeling the  $\pi^+ \pi^-$  S-wave with the K-matrix formalism. $D(2750)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D \pi$	seen
$\Gamma_2$ $D^+ \pi^-$	seen
$\Gamma_3$ $D^0 \pi^\pm$	seen
$\Gamma_4$ $D^* \pi$	seen
$\Gamma_5$ $D^{*+} \pi^-$	seen

 $D(2750)$  BRANCHING RATIOS

$$\Gamma(D^+ \pi^-)/\Gamma(D^{*+} \pi^-)$$

$$\Gamma_2/\Gamma_5$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.42 ± 0.05 ± 0.11</b>	34.8k	<sup>8</sup> DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^{*+} \pi^- X$

<sup>8</sup> The states observed in the  $D^* \pi$  and  $D \pi$  final states are not necessarily the same. $D(2750)$  POLARIZATION AMPLITUDE  $A_D$ 

A polarization amplitude  $A_D$  is a parameter that depends on the initial polarization of the  $D(2750)$ . For  $D(2750)$  decays the helicity angle,  $\theta_H$ , distribution varies like  $1 + A_D \cos(\theta_H)$ , where  $\theta_H$  is the angle in the  $D^*$  rest frame between the two pions emitted by the  $D(2750) \rightarrow D^* \pi$  and  $D^* \rightarrow D \pi$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
-0.33 ± 0.28	23.5k	<sup>9</sup> DEL-AMO-SA...10P	BABR	$e^+ e^- \rightarrow D^{*+} \pi^- X$

<sup>9</sup> Systematic uncertainties not estimated. The states observed in the  $D^* \pi$  and  $D \pi$  final states are not necessarily the same. $D(2750)$  REFERENCES

AAIJ	15Y	PR D92 032002	R. Aaij et al.	(LHCb Collab.) JP
AAIJ	13CC	JHEP 1309 145	R. Aaij et al.	(LHCb Collab.)
DEL-AMO-SA...10P	PR D82 111101	P. del Amo Sanchez et al.	(BABAR Collab.)	

 $D(3000)^0$ 

$$I(J^P) = \frac{1}{2}(?)$$

OMITTED FROM SUMMARY TABLE

Both natural- and unnatural-parity components observed depending on the decay mode (AAIJ 13CC).

 $D(3000)^0$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
2971.8 ± 8.7	9.5k	<sup>1,2</sup> AAIJ	13CC	LHCB $pp \rightarrow D^{*+} \pi^- X$
3008.1 ± 4.0	17.6k	<sup>1,3</sup> AAIJ	13CC	LHCB $pp \rightarrow D^+ \pi^- X$

<sup>1</sup> Systematic uncertainty not estimated.<sup>2</sup> Unnatural parity preferred.<sup>3</sup> Natural parity state. A state  $D(3000)^+$  is possibly seen in  $D^0 \pi^+$  final state. $D(3000)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
188.1 ± 44.8	9.5k	<sup>4,5</sup> AAIJ	13CC	LHCB $pp \rightarrow D^{*+} \pi^- X$
110.5 ± 11.5	17.6k	<sup>4,6</sup> AAIJ	13CC	LHCB $pp \rightarrow D^+ \pi^- X$

<sup>4</sup> Systematic uncertainty not estimated.<sup>5</sup> Unnatural parity preferred.<sup>6</sup> Natural parity state. A state  $D(3000)^+$  is possibly seen in  $D^0 \pi^+$  final state. $D(3000)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D^{*+} \pi^-$	seen

 $D(3000)^0$  POLARIZATION AMPLITUDE  $A_{D_J}$ 

A polarization amplitude  $A_{D_J}$  is a parameter that depends on the initial polarization of the  $D_J$ . For  $D_J$  decays the helicity angle,  $\theta_H$ , distribution varies like  $1 + A_{D_J} \cos^2(\theta_H)$ , where  $\theta_H$  is the angle in the  $D_J$  rest frame between the two pions emitted in the  $D_J \rightarrow D^* \pi$  and  $D^* \rightarrow D \pi$  decays.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
1.5 ± 0.9	9.5k	<sup>7</sup> AAIJ	13CC	LHCB $pp \rightarrow D^{*+} \pi^- X$

<sup>7</sup> Systematic uncertainty not estimated. $D(3000)^0$  REFERENCES

AAIJ	13CC	JHEP 1309 145	R. Aaij et al.	(LHCb Collab.)
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## Meson Particle Listings

 $D_s^\pm$ 

# CHARMED, STRANGE MESONS ( $C = S = \pm 1$ )

 $D_s^+ = c\bar{s}, D_s^- = \bar{c}s$ , similarly for  $D_s^{* \pm}$ 's
 $D_s^\pm$ 

$$J(P) = 0(0^-)$$

The angular distributions of the decays of the  $\phi$  and  $\bar{K}^*(892)^0$  in the  $\phi\pi^+$  and  $K^+\bar{K}^*(892)^0$  modes strongly indicate that the spin is zero. The parity given is that expected of a  $c\bar{s}$  ground state.

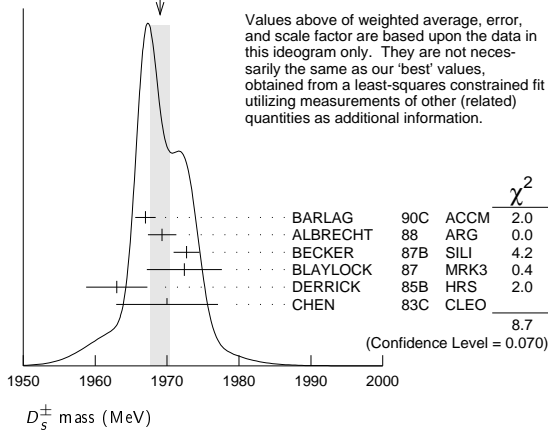
## $D_s^\pm$ MASS

The fit includes  $D_s^\pm, D^0, D_s^{* \pm}, D^{*0}, D_s^{* \pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements. Measurements of the  $D_s^\pm$  mass with an error greater than 10 MeV are omitted from the fit and average. A number of early measurements have been omitted altogether.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1968.27 ± 0.10 OUR FIT</b>				
<b>1969.0 ± 1.4 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
1967.0 ± 1.0 ± 1.0	54	BARLAG	90c	ACCM $\pi^-$ Cu 230 GeV
1969.3 ± 1.4 ± 1.4		ALBRECHT	88	ARG $e^+e^-$ 9.4–10.6 GeV
1972.7 ± 1.5 ± 1.0	21	BECKER	87b	SILI 200 GeV $\pi, K, p$
1972.4 ± 3.7 ± 3.7	27	BLAYLOCK	87	MRK3 $e^+e^-$ 4.14 GeV
1963 ± 3 ± 3	30	DERRICK	85b	HRS $e^+e^-$ 29 GeV
1970 ± 5 ± 5	104	CHEN	83c	CLEO $e^+e^-$ 10.5 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1968.3 ± 0.7 ± 0.7	290	<sup>1</sup> ANJOS	88	E691 Photoproduction
1980 ± 15	6	USHIDA	86	EMUL $\nu$ wideband
1973.6 ± 2.6 ± 3.0	163	ALBRECHT	85d	ARG $e^+e^-$ 10 GeV
1948 ± 28 ± 10	65	AIHARA	84d	TPC $e^+e^-$ 29 GeV
1975 ± 9 ± 10	49	ALTHOFF	84	TASS $e^+e^-$ 14–25 GeV
1975 ± 4	3	BAILEY	84	ACCM hadron <sup>+</sup> Be → $\phi\pi^+X$

<sup>1</sup> ANJOS 88 enters the fit via  $m_{D_s^\pm} - m_{D^\pm}$  (see below).

WEIGHTED AVERAGE  
1969.0 ± 1.4 (Error scaled by 1.5)



$$m_{D_s^\pm} - m_{D^\pm}$$

The fit includes  $D_s^\pm, D^0, D_s^{* \pm}, D^{*0}, D_s^{* \pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>98.69 ± 0.05 OUR FIT</b>				
<b>98.69 ± 0.05 OUR AVERAGE</b>				
98.68 ± 0.03 ± 0.04		AAIJ	13v	LHCB $D_s^+ \rightarrow K^+ K^- \pi^+$
99.41 ± 0.38 ± 0.21		ACOSTA	03d	CDF2 $\bar{p}p, \sqrt{s} = 1.96$ TeV
98.4 ± 0.1 ± 0.3	48k	AUBERT	02g	BABR $e^+e^- \approx \gamma(4S)$
99.5 ± 0.6 ± 0.3		BROWN	94	CLE2 $e^+e^- \approx \gamma(4S)$
98.5 ± 1.5	555	CHEN	89	CLEO $e^+e^-$ 10.5 GeV
99.0 ± 0.8	290	ANJOS	88	E691 Photoproduction

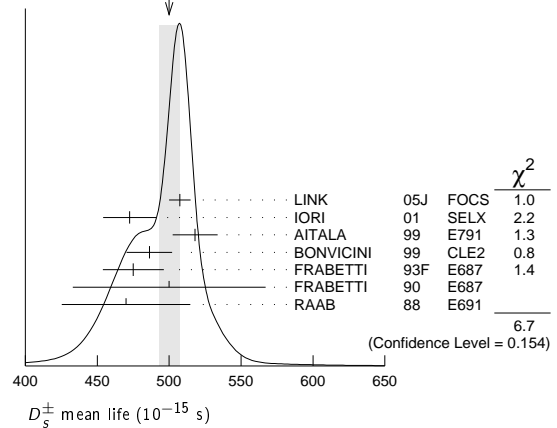
## $D_s^\pm$ MEAN LIFE

Measurements with an error greater than  $100 \times 10^{-15}$  s or with fewer than 100 events have been omitted from the Listings.

VALUE ( $10^{-15}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>500 ± 7 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
507.4 ± 5.5 ± 5.1	13.6k	LINK	05j	FOCS $\phi\pi^+$ and $\bar{K}^{*0}K^+$
472.5 ± 17.2 ± 6.6	760	IORI	01	SELX 600 GeV $\Sigma^-, \pi^-, p$
518 ± 14 ± 7	1662	AITALA	99	E791 $\pi^-$ nucleus, 500 GeV
486.3 ± 15.0 ± 4.9	2167	<sup>1</sup> BONVICINI	99	CLE2 $e^+e^- \approx \gamma(4S)$
475 ± 20 ± 7	900	FRABETTI	93f	E687 $\gamma$ Be, $\phi\pi^+$
500 ± 60 ± 30	104	FRABETTI	90	E687 $\gamma$ Be, $\phi\pi^+$
470 ± 40 ± 20	228	RAAB	88	E691 Photoproduction

<sup>1</sup> BONVICINI 99 obtains  $1.19 \pm 0.04$  for the ratio of  $D_s^+$  to  $D^0$  lifetimes.

WEIGHTED AVERAGE  
500 ± 7 (Error scaled by 1.3)



## $D_s^\pm$ DECAY MODES

Unless otherwise noted, the branching fractions for modes with a resonance in the final state include all the decay modes of the resonance.  $D_s^\pm$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor / Confidence level
<b>Inclusive modes</b>		
$\Gamma_1$ $e^+$ semileptonic	[a] ( 6.5 ± 0.4 ) %	
$\Gamma_2$ $\pi^+$ anything	( 119.3 ± 1.4 ) %	
$\Gamma_3$ $\pi^-$ anything	( 43.2 ± 0.9 ) %	
$\Gamma_4$ $\pi^0$ anything	( 123 ± 7 ) %	
$\Gamma_5$ $K^-$ anything	( 18.7 ± 0.5 ) %	
$\Gamma_6$ $K^+$ anything	( 28.9 ± 0.7 ) %	
$\Gamma_7$ $K_S^0$ anything	( 19.0 ± 1.1 ) %	
$\Gamma_8$ $\eta$ anything	[b] ( 29.9 ± 2.8 ) %	
$\Gamma_9$ $\omega$ anything	( 6.1 ± 1.4 ) %	
$\Gamma_{10}$ $\eta'$ anything	[c] ( 10.3 ± 1.4 ) %	S=1.1
$\Gamma_{11}$ $f_0(980)$ anything, $f_0 \rightarrow \pi^+\pi^-$	< 1.3 %	CL=90%
$\Gamma_{12}$ $\phi$ anything	( 15.7 ± 1.0 ) %	
$\Gamma_{13}$ $K^+K^-$ anything	( 15.8 ± 0.7 ) %	
$\Gamma_{14}$ $K_S^0K^+$ anything	( 5.8 ± 0.5 ) %	
$\Gamma_{15}$ $K_S^0K^-$ anything	( 1.9 ± 0.4 ) %	
$\Gamma_{16}$ $2K_S^0$ anything	( 1.70 ± 0.32 ) %	
$\Gamma_{17}$ $2K^+$ anything	< 2.6 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{18}$ $2K^-$ anything	< 6 × 10 <sup>-4</sup>	CL=90%
<b>Leptonic and semileptonic modes</b>		
$\Gamma_{19}$ $e^+\nu_e$	< 8.3 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{20}$ $\mu^+\nu_\mu$	( 5.56 ± 0.25 ) × 10 <sup>-3</sup>	
$\Gamma_{21}$ $\tau^+\nu_\tau$	( 5.55 ± 0.24 ) %	
$\Gamma_{22}$ $K^+K^-e^+\nu_e$	—	
$\Gamma_{23}$ $\phi e^+\nu_e$	[d] ( 2.39 ± 0.23 ) %	S=1.8
$\Gamma_{24}$ $\eta e^+\nu_e + \eta'(958) e^+\nu_e$	[d] ( 2.96 ± 0.29 ) %	
$\Gamma_{25}$ $\eta e^+\nu_e$	[d] ( 2.28 ± 0.24 ) %	
$\Gamma_{26}$ $\eta'(958) e^+\nu_e$	[d] ( 6.8 ± 1.6 ) × 10 <sup>-3</sup>	
$\Gamma_{27}$ $\omega e^+\nu_e$	[e] < 2.0 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{28}$ $K^0 e^+\nu_e$	( 3.9 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{29}$ $K^*(892)^0 e^+\nu_e$	[d] ( 1.8 ± 0.4 ) × 10 <sup>-3</sup>	
$\Gamma_{30}$ $f_0(980) e^+\nu_e, f_0 \rightarrow \pi^+\pi^-$		



Hadronic modes with a $K\bar{K}$ pair			
$\Gamma_{31}$	$K^+ K_S^0$	( 1.50±0.05 ) %	S=1.2
$\Gamma_{32}$	$K^+ \bar{K}^0$	( 2.95±0.14 ) %	
$\Gamma_{33}$	$K^+ K^- \pi^+$	[f] ( 5.45±0.17 ) %	
$\Gamma_{34}$	$\phi \pi^+$	[d,g] ( 4.5 ±0.4 ) %	
$\Gamma_{35}$	$\phi \pi^+, \phi \rightarrow K^+ K^-$	[g] ( 2.27±0.08 ) %	CL=90%
$\Gamma_{36}$	$K^+ \bar{K}^*(892)^0, \bar{K}^{*0} \rightarrow$	( 2.61±0.09 ) %	
$\Gamma_{37}$	$K^- \pi^+$	( 1.15±0.32 ) %	
$\Gamma_{38}$	$f_0(980) \pi^+, f_0 \rightarrow K^+ K^-$	( 7 ±5 ) ×10 <sup>-4</sup>	
$\Gamma_{39}$	$f_0(1370) \pi^+, f_0 \rightarrow K^+ K^-$	( 6.7 ±2.9 ) ×10 <sup>-4</sup>	CL=90%
$\Gamma_{40}$	$f_0(1710) \pi^+, f_0 \rightarrow K^+ K^-$	( 6.7 ±2.9 ) ×10 <sup>-4</sup>	
$\Gamma_{41}$	$K^+ \bar{K}_0^*(1430)^0, \bar{K}_0^{*0} \rightarrow$	( 1.9 ±0.4 ) ×10 <sup>-3</sup>	
$\Gamma_{42}$	$K^- \pi^+$	( 1.52±0.22 ) %	

$\Gamma_{43}$	$2K_S^0 \pi^+$	( 7.7 ±0.6 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{44}$	$K^0 \bar{K}^0 \pi^+$	—	
$\Gamma_{45}$	$K^*(892)^+ \bar{K}^0$	[d] ( 5.4 ±1.2 ) %	
$\Gamma_{46}$	$K^+ K^- \pi^+ \pi^0$	( 6.3 ±0.6 ) %	
$\Gamma_{47}$	$\phi \rho^+$	[d] ( 8.4 $^{+1.9}_{-2.3}$ ) %	CL=90%
$\Gamma_{48}$	$K_S^0 K^- 2\pi^+$	( 1.67±0.10 ) %	
$\Gamma_{49}$	$K^*(892)^+ \bar{K}^*(892)^0$	[d] ( 7.2 ±2.6 ) %	
$\Gamma_{50}$	$K^+ K_S^0 \pi^+ \pi^-$	( 1.03±0.10 ) %	
$\Gamma_{51}$	$K^+ K^- 2\pi^+ \pi^-$	( 8.7 ±1.5 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{52}$	$\phi 2\pi^+ \pi^-$	[d] ( 1.21±0.16 ) %	
$\Gamma_{53}$	$K^+ K^- \rho^0 \pi^+ \text{non-}\phi$	< 2.6 ×10 <sup>-4</sup>	
$\Gamma_{54}$	$\phi \rho^0 \pi^+, \phi \rightarrow K^+ K^-$	( 6.5 ±1.3 ) ×10 <sup>-3</sup>	
$\Gamma_{55}$	$\phi a_1(1260)^+, \phi \rightarrow$	( 7.5 ±1.2 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{56}$	$K^+ K^-, a_1^+ \rightarrow \rho^0 \pi^+$	( 9 ±7 ) ×10 <sup>-4</sup>	
$\Gamma_{57}$	$K^+ K^- 2\pi^+ \pi^- \text{nonresonant}$	( 9 ±4 ) ×10 <sup>-4</sup>	
$\Gamma_{58}$	$2K_S^0 2\pi^+ \pi^-$	( 9 ±4 ) ×10 <sup>-4</sup>	

Hadronic modes without $K$ 's			
$\Gamma_{57}$	$\pi^+ \pi^0$	< 3.5 ×10 <sup>-4</sup>	CL=90%
$\Gamma_{58}$	$2\pi^+ \pi^-$	( 1.09±0.05 ) %	
$\Gamma_{59}$	$\rho^0 \pi^+$	( 2.0 ±1.2 ) ×10 <sup>-4</sup>	
$\Gamma_{60}$	$\pi^+ (\pi^+ \pi^-)_{S\text{-wave}}$	[h] ( 9.1 ±0.4 ) ×10 <sup>-3</sup>	
$\Gamma_{61}$	$f_0(980) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	( 1.10±0.20 ) ×10 <sup>-3</sup>	S=1.1
$\Gamma_{62}$	$f_0(1370) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	( 3.0 ±2.0 ) ×10 <sup>-4</sup>	
$\Gamma_{63}$	$f_0(1500) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	( 6.5 ±1.3 ) ×10 <sup>-3</sup>	
$\Gamma_{64}$	$f_2(1270) \pi^+, f_2 \rightarrow \pi^+ \pi^-$	( 1.10±0.20 ) ×10 <sup>-3</sup>	
$\Gamma_{65}$	$\rho(1450)^0 \pi^+, \rho^0 \rightarrow \pi^+ \pi^-$	( 3.0 ±2.0 ) ×10 <sup>-4</sup>	CL=90%
$\Gamma_{66}$	$\pi^+ 2\pi^0$	( 6.5 ±1.3 ) ×10 <sup>-3</sup>	
$\Gamma_{67}$	$2\pi^+ \pi^- \pi^0$	—	
$\Gamma_{68}$	$\eta \pi^+$	[d] ( 1.70±0.09 ) %	
$\Gamma_{69}$	$\omega \pi^+$	[d] ( 2.4 ±0.6 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{70}$	$3\pi^+ 2\pi^-$	( 8.0 ±0.8 ) ×10 <sup>-3</sup>	
$\Gamma_{71}$	$2\pi^+ \pi^- 2\pi^0$	—	
$\Gamma_{72}$	$\eta \rho^+$	[d] ( 8.9 ±0.8 ) %	
$\Gamma_{73}$	$\eta \pi^+ \pi^0$	( 9.2 ±1.2 ) %	CL=90%
$\Gamma_{74}$	$\omega \pi^+ \pi^0$	[d] ( 2.8 ±0.7 ) %	
$\Gamma_{75}$	$3\pi^+ 2\pi^- \pi^0$	( 4.9 ±3.2 ) %	
$\Gamma_{76}$	$\omega 2\pi^+ \pi^-$	[d] ( 1.6 ±0.5 ) %	
$\Gamma_{77}$	$\eta'(958) \pi^+$	[c,d] ( 3.94±0.25 ) %	CL=90%
$\Gamma_{78}$	$3\pi^+ 2\pi^- 2\pi^0$	—	
$\Gamma_{79}$	$\omega \eta \pi^+$	[d] < 2.13 %	
$\Gamma_{80}$	$\eta'(958) \rho^+$	[c,d] ( 5.8 ±1.5 ) %	
$\Gamma_{81}$	$\eta'(958) \pi^+ \pi^0$	( 5.6 ±0.8 ) %	CL=90%
$\Gamma_{82}$	$\eta'(958) \pi^+ \pi^0 \text{nonresonant}$	< 5.1 %	

Modes with one or three $K$ 's			
$\Gamma_{83}$	$K^+ \pi^0$	( 6.3 ±2.1 ) ×10 <sup>-4</sup>	CL=90%
$\Gamma_{84}$	$K_S^0 \pi^+$	( 1.22±0.06 ) ×10 <sup>-3</sup>	
$\Gamma_{85}$	$K^+ \eta$	[d] ( 1.77±0.35 ) ×10 <sup>-3</sup>	
$\Gamma_{86}$	$K^+ \omega$	[d] < 2.4 ×10 <sup>-3</sup>	
$\Gamma_{87}$	$K^+ \eta'(958)$	[d] ( 1.8 ±0.6 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{88}$	$K^+ \pi^+ \pi^-$	( 6.6 ±0.4 ) ×10 <sup>-3</sup>	
$\Gamma_{89}$	$K^+ \rho^0$	( 2.5 ±0.4 ) ×10 <sup>-3</sup>	
$\Gamma_{90}$	$K^+ \rho(1450)^0, \rho^0 \rightarrow \pi^+ \pi^-$	( 7.0 ±2.4 ) ×10 <sup>-4</sup>	
$\Gamma_{91}$	$K^*(892)^0 \pi^+, K^{*0} \rightarrow K^+ \pi^-$	( 1.42±0.24 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{92}$	$K^*(1410)^0 \pi^+, K^{*0} \rightarrow$	( 1.24±0.29 ) ×10 <sup>-3</sup>	
$\Gamma_{93}$	$K^+ \pi^-$	( 5.0 ±3.5 ) ×10 <sup>-4</sup>	
$\Gamma_{94}$	$K^*(1430)^0 \pi^+, K^{*0} \rightarrow$	( 1.04±0.34 ) ×10 <sup>-3</sup>	
$\Gamma_{95}$	$K^+ \pi^+ \pi^- \text{nonresonant}$	( 1.00±0.18 ) %	

$\Gamma_{96}$	$K_S^0 2\pi^+ \pi^-$	( 3.0 ±1.1 ) ×10 <sup>-3</sup>	CL=90%
$\Gamma_{97}$	$K^+ \omega \pi^0$	[d] < 8.2 ×10 <sup>-3</sup>	
$\Gamma_{98}$	$K^+ \omega \pi^+ \pi^-$	[d] < 5.4 ×10 <sup>-3</sup>	
$\Gamma_{99}$	$K^+ \omega \eta$	[d] < 7.9 ×10 <sup>-3</sup>	
$\Gamma_{100}$	$2K^+ K^-$	( 2.18±0.21 ) ×10 <sup>-4</sup>	CL=90%
$\Gamma_{101}$	$\phi K^+, \phi \rightarrow K^+ K^-$	( 8.9 ±2.0 ) ×10 <sup>-5</sup>	

## Doubly Cabibbo-suppressed modes

$\Gamma_{102}$	$2K^+ \pi^-$	( 1.27±0.13 ) ×10 <sup>-4</sup>
$\Gamma_{103}$	$K^+ K^*(892)^0, K^{*0} \rightarrow$	( 6.0 ±3.4 ) ×10 <sup>-5</sup>
	$K^+ \pi^-$	

## Baryon-antibaryon mode

$\Gamma_{104}$	$p \bar{n}$	( 1.3 ±0.4 ) ×10 <sup>-3</sup>
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 $\Delta C = 1$  weak neutral current ( $C1$ ) modes,  
Lepton family number ( $LF$ ), or  
Lepton number ( $L$ ) violating modes

$\Gamma_{105}$	$\pi^+ e^+ e^-$	[l] < 1.3 ×10 <sup>-5</sup>	CL=90%
$\Gamma_{106}$	$\pi^+ \phi, \phi \rightarrow e^+ e^-$	[l] ( 6 $^{+8}_{-4}$ ) ×10 <sup>-6</sup>	
$\Gamma_{107}$	$\pi^+ \mu^+ \mu^-$	[l] < 4.1 ×10 <sup>-7</sup>	
$\Gamma_{108}$	$K^+ e^+ e^-$	$C1$ < 3.7 ×10 <sup>-6</sup>	
$\Gamma_{109}$	$K^+ \mu^+ \mu^-$	$C1$ < 2.1 ×10 <sup>-5</sup>	CL=90%
$\Gamma_{110}$	$K^*(892)^+ \mu^+ \mu^-$	$C1$ < 1.4 ×10 <sup>-3</sup>	
$\Gamma_{111}$	$\pi^+ e^+ \mu^-$	$LF$ < 1.2 ×10 <sup>-5</sup>	
$\Gamma_{112}$	$\pi^+ e^- \mu^+$	$LF$ < 2.0 ×10 <sup>-5</sup>	
$\Gamma_{113}$	$K^+ e^+ \mu^-$	$LF$ < 1.4 ×10 <sup>-5</sup>	CL=90%
$\Gamma_{114}$	$K^+ e^- \mu^+$	$LF$ < 9.7 ×10 <sup>-6</sup>	
$\Gamma_{115}$	$\pi^- 2e^+$	$L$ < 4.1 ×10 <sup>-6</sup>	
$\Gamma_{116}$	$\pi^- 2\mu^+$	$L$ < 1.2 ×10 <sup>-7</sup>	
$\Gamma_{117}$	$\pi^- e^+ \mu^+$	$L$ < 8.4 ×10 <sup>-6</sup>	CL=90%
$\Gamma_{118}$	$K^- 2e^+$	$L$ < 5.2 ×10 <sup>-6</sup>	
$\Gamma_{119}$	$K^- 2\mu^+$	$L$ < 1.3 ×10 <sup>-5</sup>	
$\Gamma_{120}$	$K^- e^+ \mu^+$	$L$ < 6.1 ×10 <sup>-6</sup>	
$\Gamma_{121}$	$K^*(892)^- 2\mu^+$	$L$ < 1.4 ×10 <sup>-3</sup>	CL=90%

[a] This is the purely  $e^+$  semileptonic branching fraction: the  $e^+$  fraction from  $\tau^+$  decays has been subtracted off. The sum of our (non- $\tau$ )  $e^+$  exclusive fractions — an  $e^+ \nu_e$  with an  $\eta, \eta', \phi, K^0, K^{*0}$ , or  $f_0(980)$  — is  $7.0 \pm 0.4$  %

[b] This fraction includes  $\eta$  from  $\eta'$  decays.

[c] Two times (to include  $\mu$  decays) the  $\eta' e^+ \nu_e$  branching fraction, plus the  $\eta' \pi^+, \eta' \rho^+$ , and  $\eta' K^+$  fractions, is  $(18.6 \pm 2.3)\%$ , which considerably exceeds the inclusive  $\eta'$  fraction of  $(11.7 \pm 1.8)\%$ . Our best guess is that the  $\eta' \rho^+$  fraction,  $(12.5 \pm 2.2)\%$ , is too large.

[d] This branching fraction includes all the decay modes of the final-state resonance.

[e] A test for  $u\bar{u}$  or  $d\bar{d}$  content in the  $D_s^+$ . Neither Cabibbo-favored nor Cabibbo-suppressed decays can contribute, and  $\omega$ - $\phi$  mixing is an unlikely explanation for any fraction above about  $2 \times 10^{-4}$ .

[f] The branching fraction for this mode may differ from the sum of the submodes that contribute to it, due to interference effects. See the relevant papers.

[g] We decouple the  $D_s^+ \rightarrow \phi \pi^+$  branching fraction obtained from mass projections (and used to get some of the other branching fractions) from the  $D_s^+ \rightarrow \phi \pi^+, \phi \rightarrow K^+ K^-$  branching fraction obtained from the Dalitz-plot analysis of  $D_s^+ \rightarrow K^+ K^- \pi^+$ . That is, the ratio of these two branching fractions is not exactly the  $\phi \rightarrow K^+ K^-$  branching fraction 0.491.

[h] This is the average of a model-independent and a  $K$ -matrix parametrization of the  $\pi^+ \pi^-$   $S$ -wave and is a sum over several  $f_0$  mesons.

[i] This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.

[j] This is *not* a test for the  $\Delta C=1$  weak neutral current, but leads to the  $\pi^+ \ell^+ \ell^-$  final state.

# Meson Particle Listings

 $D_s^\pm$ 

## CONSTRAINED FIT INFORMATION

An overall fit to 14 branching ratios uses 18 measurements and one constraint to determine 12 parameters. The overall fit has a  $\chi^2 = 8.1$  for 7 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_{25}$	0												
$x_{26}$	0	0											
$x_{31}$	0	0	0										
$x_{33}$	0	0	0	56									
$x_{45}$	0	0	0	15	27								
$x_{47}$	0	0	0	35	34	11							
$x_{58}$	0	0	0	36	55	16	22						
$x_{68}$	0	0	0	16	1	-2	7	-1					
$x_{69}$	0	0	0	2	0	0	1	0	11				
$x_{88}$	0	0	0	21	20	3	12	10	11	1			
	$x_{23}$	$x_{25}$	$x_{26}$	$x_{31}$	$x_{33}$	$x_{45}$	$x_{47}$	$x_{58}$	$x_{68}$	$x_{69}$			

## $D_s^+$ BRANCHING FRACTIONS

Updated November 2015 by J.L. Rosner (University of Chicago) and C.G. Wohl (LBNL).

Figure 1 shows a partial breakdown of the  $D_s^+$  branching fractions. The rest of this note is about how the figure was constructed. The values shown make heavy use of CLEO measurements of inclusive branching fractions [1]. For references to other data cited in the following, see the Listings.

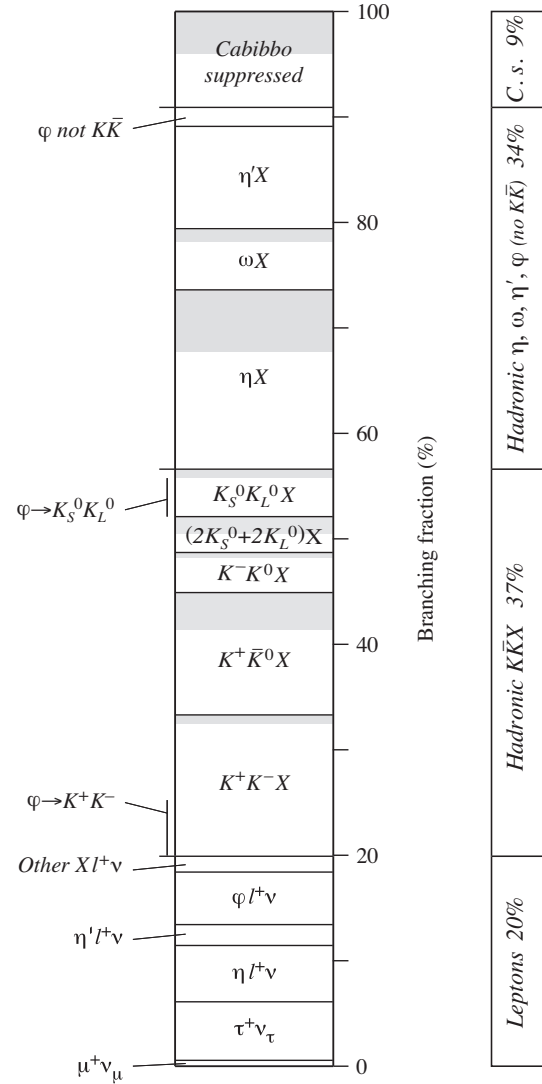
**Modes with leptons:** The bottom  $(19.9 \pm 0.9)\%$  of Fig. 1 shows the fractions for the modes that include leptons. Measured  $Xe^+\nu_e$  semileptonic fractions have been doubled to include the  $X\mu^+\nu_\mu$  fractions. The sum of the exclusive  $Xe^+\nu_e$  fractions is  $(6.9 \pm 0.4)\%$ , consistent with an inclusive semileptonic measurement of  $(6.5 \pm 0.4)\%$ . There seems to be little missing here.

**Inclusive hadronic  $K\bar{K}$  fractions:** The Cabibbo-favored  $c \rightarrow s$  decay in  $D_s^+$  decay produces a final state with both an  $s$  and an  $\bar{s}$ ; and thus modes with a  $K\bar{K}$  pair or with an  $\eta$ ,  $\omega$ ,  $\eta'$ , or  $\phi$  predominate (as may already be seen in Fig. 1 in the semileptonic fractions). We consider the  $K\bar{K}$  modes first. A complete picture of the exclusive  $K\bar{K}$  charge modes is not yet possible, because branching fractions for many of those modes have not yet been measured. However, CLEO has measured the inclusive  $K^+$ ,  $K^-$ ,  $K_S^0$ ,  $K^+K^-$ ,  $K^+K_S^0$ ,  $K^-K_S^0$ , and  $2K_S^0$  fractions (these include modes with leptons) [1]. And each of these inclusive fractions with a  $K_S^0$  is equal to the corresponding fraction with a  $K_L^0$ :  $f(K^+K_L^0) = f(K^+K_S^0)$ ,  $f(2K_L^0) = f(2K_S^0)$ , *etc.* Therefore, of all inclusive fractions pairing a  $K^+$ ,  $K_S^0$ , or  $K_L^0$  with a  $K^-$ ,  $K_S^0$ , or  $K_L^0$ , we know all but  $f(K_S^0K_L^0)$ .

We can get that fraction. The total  $K_S^0$  fraction is

$$f(K_S^0) = f(K^+K_S^0) + f(K^-K_S^0) + 2f(2K_S^0) + f(K_S^0K_L^0) + f(\text{single } K_S^0),$$

where  $f(\text{single } K_S^0)$  is the sum of the branching fractions for modes such as  $K_S^0\pi^+2\pi^0$  with a  $K_S^0$  and no second  $K$ . The  $K_S^0\pi^+2\pi^0$  mode is in fact the only unmeasured single- $K_S^0$  mode



**Figure 1:** A partial breakdown of  $D_s^+$  branching fractions. The hadronic bins in the left column show inclusive fractions. Shading within a bin shows how much of the inclusive fraction is not yet accounted for by adding up all the relevant exclusive fractions. The inclusive hadronic  $\phi$  fraction is spread over three bins, in proportion to its decay fractions into  $K^+K^-$ ,  $K_S^0K_L^0$ , and no- $K\bar{K}$  modes.

(throughout, we shall assume that fractions for modes with a  $K$  or  $K\bar{K}$  and more than three pions are negligible), and we shall take its fraction to be the same as for the  $K_S^0 2\pi^+\pi^-$  mode,  $(0.30 \pm 0.11)\%$ . Any reasonable deviation from this value would be too small to matter much in the following. Adding the several small single- $K_S^0$  branching fractions, including those from semileptonic modes, we get  $f(\text{single } K_S^0) = (1.65 \pm 0.26)\%$ .

Using this, we have:

$$\begin{aligned} f(K_S^0K_L^0) &= f(K_S^0) - f(K^+K_S^0) - f(K^-K_S^0) \\ &\quad - 2f(2K_S^0) - f(\text{single } K_S^0) \\ &= (19.0 \pm 1.1) - (5.8 \pm 0.5) - (1.9 \pm 0.4) \end{aligned}$$

$$\begin{aligned}
& -2 \times (1.7 \pm 0.3) - (1.7 \pm 0.3) \\
& = (6.2 \pm 1.4)\% .
\end{aligned}$$

Here and below we treat the errors as uncorrelated, although often they are not. However, our main aim is to get numbers for Fig. 1; errors are secondary.

There is a check on our result: The  $\phi$  inclusive branching fraction is  $(15.7 \pm 1.0)\%$ , of which 34%, or  $(5.34 \pm 0.34)\%$  of  $D_s^+$  decays, produces a  $K_S^0 K_L^0$ . Our  $f(K_S^0 K_L^0) = (6.2 \pm 1.4)\%$  has to be at least this large—and it is.

We now have all the inclusive  $K\bar{K}$  fractions. We use  $f(K^+ \bar{K}^0) = 2 f(K^+ K_S^0)$ , and likewise for  $f(K^- K^0)$ . For  $K^+ K^-$  and  $K_S^0 K_L^0$ , we subtract off the contributions from  $\phi \ell^+ \nu$  decay to get the purely hadronic  $K\bar{K}$  inclusive fractions:

$$\begin{aligned}
f(K^+ K^-, \text{hadronic}) &= (15.8 \pm 0.7) - (2.44 \pm 0.14) \\
&= (13.4 \pm 0.7)\% \\
f(K^+ \bar{K}^0, \text{hadronic}) &= (11.6 \pm 1.0)\% \\
f(K^- K^0, \text{hadronic}) &= (3.8 \pm 0.8)\% \\
f(2K_S^0 + 2K_L^0, \text{hadronic}) &= (3.4 \pm 0.64)\% \\
f(K_S^0 K_L^0, \text{hadronic}) &= (6.2 \pm 1.4) - (1.70 \pm 0.10) \\
&= (4.5 \pm 1.4)\% .
\end{aligned}$$

The fractions are shown in Fig. 1. They total  $(36.7 \pm 2.1)\%$  of  $D_s^+$  decays.

We can add more information to the figure by summing up measured branching fractions for exclusive modes within each bin:

**$K^+ K^-$  modes**—The sum of measured  $K^+ K^- \pi^+$ ,  $K^+ K^- \pi^+ \pi^0$ , and  $K^+ K^- 2\pi^+ \pi^-$  branching fractions is  $(12.6 \pm 0.6)\%$ . That leaves  $(0.8 \pm 0.9)\%$  for the  $K^+ K^- \pi^+ 2\pi^0$  mode, which is the only other  $K^+ K^-$  mode with three or fewer pions. In Fig. 1, this unmeasured part of the  $K^+ K^-$  bin is shaded.

**$K^+ \bar{K}^0$  modes**—Two times the sum of the measured  $K^+ K_S^0$ ,  $K^+ K_S^0 \pi^0$ , and  $K^+ K_S^0 \pi^+ \pi^-$  branching fractions is  $(8.1 \pm 0.5)\%$ . This leaves  $(3.5 \pm 1.1)\%$  for the unmeasured  $K^+ \bar{K}^0$  modes (there are three such modes with three or fewer pions). This is shaded in the figure.

**$K^- K^0$  modes**—Twice the  $K^- K_S^0 2\pi^+$  fraction is  $(3.34 \pm 0.20)\%$ , which leaves about  $(0.5 \pm 0.8)\%$  for  $K^- K^0 2\pi^+ \pi^0$ , the only other  $K^- K^0$  mode with three or fewer pions.

**$2K_S^0 + 2K_L^0$  modes**—The  $2K_S^0 \pi^+$  and  $2K_S^0 2\pi^+ \pi^-$  fractions sum to  $(0.86 \pm 0.07)\%$ ; this times two (for the corresponding  $2K_L^0$  modes) is  $(1.72 \pm 0.14)\%$ . This leaves about  $(1.7 \pm 0.7)\%$  for other  $2K_S^0 + 2K_L^0$  modes.

**$K_S^0 K_L^0$  modes**—Most of the  $K_S^0 K_L^0$  fraction is accounted for by  $\phi$  decays (see below).

**Inclusive hadronic  $\eta$ ,  $\omega$ ,  $\eta'$ , and  $\phi$  fractions:** These are easier. We start with the inclusive branching fractions, and then, to avoid double counting, subtract: (1) fractions for modes

with leptons; (2)  $\eta$  mesons that are included in the inclusive  $\eta'$  fraction; and (3)  $K^+ K^-$  and  $K_S^0 K_L^0$  from  $\phi$  decays:

$$\begin{aligned}
f(\eta \text{ hadronic}) &= f(\eta \text{ inclusive}) - 0.65 f(\eta' \text{ inclusive}) \\
&\quad - f(\eta \ell^+ \nu) = (17.0 \pm 3.1)\% \\
f(\omega \text{ hadronic}) &= f(\omega \text{ inclusive}) - 0.0275 f(\eta' \text{ inclusive}) \\
&\quad = (5.8 \pm 1.4)\% \\
f(\eta' \text{ hadronic}) &= f(\eta' \text{ inclusive}) - f(\eta' \ell^+ \nu) \\
&\quad = (9.7 \pm 1.9)\% \\
f(\phi \text{ hadronic}, \not\rightarrow K\bar{K}) &= 0.17 [f(\phi \text{ inclusive}) - f(\phi \ell^+ \nu)] \\
&\quad = (1.8 \pm 0.2)\% .
\end{aligned}$$

The factors 0.65, 0.0275, and 0.17 are the  $\eta' \rightarrow \eta$ ,  $\eta' \rightarrow \omega$ , and  $\phi \rightarrow K\bar{K}$  branching fractions. Figure 1 shows the results; the sum is  $(34.2 \pm 3.9)\%$ , which is about equal to the hadronic  $K\bar{K}$  total.

Note that the bin marked  $\phi$  near the top of Fig. 1 includes neither the  $\phi \ell^+ \nu$  decays nor the 83% of other  $\phi$  decays that produce a  $K\bar{K}$  pair. There is twice as much  $\phi$  in the  $K_S^0 K_L^0$  bin, and nearly three times as much in the  $K^+ K^-$  bin. These contributions are indicated in those bins.

Again, we can show how much of each bin is accounted for by measured exclusive branching fractions:

**$\eta$  modes**—The sum of  $\eta \pi^+$ ,  $\eta \rho^+$ , and  $\eta K^+$  branching fractions is  $(11.1 \pm 1.2)\%$ , which leaves a good part of the inclusive hadronic  $\eta$  fraction,  $(17.0 \pm 3.1)\%$ , to be accounted for. This is shaded in the figure.

**$\omega$  modes**—The sum of  $\omega \pi^+$ ,  $\omega \pi^+ \pi^0$ , and  $\omega 2\pi^+ \pi^-$  fractions is  $(4.6 \pm 0.9)\%$ , which is nearly as large as the inclusive hadronic  $\omega$  fraction,  $(5.8 \pm 1.4)\%$ .

**$\eta'$  modes**—The sum of  $\eta' \pi^+$ ,  $\eta' \rho^+$ , and  $\eta' K^+$  fractions is  $(9.7 \pm 1.9)\%$ , which agrees with the inclusive hadronic  $\eta'$  fraction,  $(9.7 \pm 1.9)\%$ . (An old measurement of the  $\eta' \rho^+$  fraction,  $(12.5 \pm 2.2)\%$ , has been abandoned [2].)

**Cabibbo-suppressed modes:** The sum of the fractions for modes with a  $K\bar{K}$ ,  $\eta$ ,  $\omega$ ,  $\eta'$ , or leptons is  $(90.8 \pm 4.5)\%$ . The remaining  $(9.2 \pm 4.5)\%$  is to Cabibbo-suppressed modes, mainly single- $K$ + pions and multiple-pion modes (see below). However, it should be noted that some small parts of the modes already discussed are Cabibbo-suppressed. For example, the  $(1.10 \pm 0.24)\%$  of  $D_s^+$  decays to  $K^0 \ell \nu$  or  $K^{*0} \ell \nu$  is already in the  $X \ell \nu$  bin in Fig. 1. And the inclusive measurements of  $\eta$ ,  $\omega$ , and  $\eta'$  fractions do not distinguish between (and therefore include both) Cabibbo-allowed and -suppressed modes. We shall not try to make a separation here.

**$K^0$  + pions**—Above, we found that  $f(\text{single } K_S^0) = (1.65 \pm 0.26)\%$ . Subtracting leptonic fractions with a  $K_S^0$  leaves  $(1.22 \pm 0.28)\%$ . The hadronic single- $K^0$  fraction is twice this,  $(2.44 \pm 0.56)\%$ . The sum of measured  $K^0 \pi^+$ ,  $K^0 \pi^+ \pi^0$ , and  $K^0 2\pi^+ \pi^-$  fractions is  $(1.84 \pm 0.28)\%$ .

**$K^+ + \text{pions}$** —The  $K^+ \pi^0$  and  $K^+ \pi^+ \pi^-$  fractions sum to  $(0.72 \pm 0.05)\%$ . Much of the  $K^+ n \pi$  modes, where  $n \geq 3$ , is

# Meson Particle Listings

$D_s^\pm$

already in the  $\eta$ ,  $\omega$ , and  $\eta'$  bins, and the rest is not measured. The total  $K^+$  fraction wanted here is probably in the 1-to-2% range.

*Multi-pions*—The  $2\pi^+\pi^-$ ,  $\pi^+2\pi^0$ , and  $3\pi^+2\pi^-$  fractions total  $(2.54 \pm 0.16)\%$ . Modes not measured might double this.

The sum of the actually measured fractions is  $(5.1 \pm 0.3)\%$ , which is not inconsistent with the Cabibbo-suppressed total of  $(9.2 \pm 4.5)\%$ .

**A model:** With CLEO about to publish inclusive branching fractions [1], Gronau and Rosner predicted those fractions using a “statistical isospin” model [3]. Consider, say, the  $D_s^+ \rightarrow K\bar{K}\pi$  charge modes: the  $K^+K^-\pi^+$  branching fraction is measured, the  $K^+\bar{K}^0\pi^0$  and  $K^0\bar{K}^0\pi^+$  fractions are not. The statistical isospin model assumes that all the independent isospin amplitudes for  $D_s^+ \rightarrow K\bar{K}\pi$  decay are equal in magnitude and incoherent in phase—in which case, the ratio of the three fractions here is 3:3:2. (Actually, use was also made of the fact that  $D_s^+ \rightarrow K\bar{K}\pi$  decay is dominated by  $\phi\pi^+$ ,  $K^+\bar{K}^{*0}$ , and  $K^{*+}\bar{K}^0$  submodes; but the estimated charge-mode ratios were not far from 3:3:2.) A different, quark-antiquark pair-production model was used to estimate systematic uncertainties.

In this way, unmeasured exclusive fractions were calculated from measured exclusive fractions (the latter were taken from the 2008 Review, and so did not benefit from recent results). In the hadronic sector, the measured total of 59.4% of  $D_s^+$  decays led to an estimated total of 24.2% for unmeasured modes. Weighted counts of  $\pi^+$ ,  $K_S^0$ , *etc.*, were then made to get the inclusive fractions.

Of interest here is that the sum of all the exclusive fractions—a way-stop in getting the inclusive values—was a nearly correct 103%. In the absence of complete measurements, the model is a way to, in effect, average over ignorance. It probably works better summed over a number of charge-mode sets than in detail. It is known to sometimes give incorrect results when there are sufficient measurements to test it.

## References

1. S. Dobbs *et al.*, Phys. Rev. **D79**, 112008 (2009).
2. P.U.E. Onyisi *et al.*, Phys. Rev. **D88**, 032009 (2013).
3. M. Gronau, J.L. Rosner, Phys. Rev. **D79**, 074022 (2009).

## $D_s^\pm$ BRANCHING RATIOS

A number of older, now obsolete results have been omitted. They may be found in earlier editions.

### Inclusive modes

$\Gamma(e^+ \text{ semileptonic})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
This is the purely $e^+$ semileptonic branching fraction: the $e^+$ fraction from $\tau^+$ decays has been subtracted off. The sum of our (non- $\tau$ ) $e^+$ exclusive fractions — an $e^+\nu_e$ with an $\eta$ , $\eta'$ , $\phi$ , $K^0$ , $K^{*0}$ , or $f_0(980)$ — is $6.90 \pm 0.4\%$				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.52 \pm 0.39 \pm 0.15</math></b>	$536 \pm 29$	<sup>1</sup> ASNER	10	CLEO $e^+e^-$ at 3774 MeV

<sup>1</sup> Using the  $D_s^+$  and  $D^0$  lifetimes, ASNER 10 finds that the ratio of the  $D_s^+$  and  $D^0$  semileptonic widths is  $0.828 \pm 0.051 \pm 0.025$ .

$\Gamma(\pi^+ \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
Events with two $\pi^+$ 's count twice, etc. But $\pi^+$ 's from $K_S^0 \rightarrow \pi^+\pi^-$ are not included.				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>119.3<math>\pm</math>1.2<math>\pm</math>0.7</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(\pi^- \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
Events with two $\pi^-$ 's count twice, etc. But $\pi^-$ 's from $K_S^0 \rightarrow \pi^+\pi^-$ are not included.				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>43.2<math>\pm</math>0.9<math>\pm</math>0.3</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(\pi^0 \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
Events with two $\pi^0$ 's count twice, etc. But $\pi^0$ 's from $K_S^0 \rightarrow 2\pi^0$ are not included.				
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>123.4<math>\pm</math>3.8<math>\pm</math>5.3</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(K^- \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>18.7<math>\pm</math>0.5<math>\pm</math>0.2</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(K^+ \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>28.9<math>\pm</math>0.6<math>\pm</math>0.3</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(K_S^0 \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>19.0<math>\pm</math>1.0<math>\pm</math>0.4</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(\eta \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
This ratio includes $\eta$ particles from $\eta'$ decays.				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>29.9<math>\pm</math>2.2<math>\pm</math>1.7</b>		DOBBS	09	CLEO $e^+e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
23.5 $\pm$ 3.1 $\pm$ 2.0	674 $\pm$ 91	HUANG	06B	CLEO See DOBBS 09
$\Gamma(\omega \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>6.1<math>\pm</math>1.4<math>\pm</math>0.3</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(\eta' \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.3<math>\pm</math>1.4 OUR AVERAGE</b>	Error includes scale factor of 1.1.			
8.8 $\pm$ 1.8 $\pm$ 0.5	68	ABLIKIM	15Z	BES3 482 pb <sup>-1</sup> , 4009 MeV
11.7 $\pm$ 1.7 $\pm$ 0.7		DOBBS	09	CLEO $e^+e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.7 $\pm$ 1.9 $\pm$ 0.8	68	HUANG	06B	CLEO See DOBBS 09
$\Gamma(f_0(980) \text{ anything, } f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.3</b>	90	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV
$\Gamma(\phi \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>15.7<math>\pm</math>0.8<math>\pm</math>0.6</b>		DOBBS	09	CLEO $e^+e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
16.1 $\pm$ 1.2 $\pm$ 1.1	398 $\pm$ 27	HUANG	06B	CLEO See DOBBS 09
$\Gamma(K^+K^- \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>15.8<math>\pm</math>0.6<math>\pm</math>0.3</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(K_S^0K^+ \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>5.8<math>\pm</math>0.5<math>\pm</math>0.1</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(K_S^0K^- \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{15}/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.9<math>\pm</math>0.4<math>\pm</math>0.1</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(2K_S^0 \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{16}/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.7<math>\pm</math>0.3<math>\pm</math>0.1</b>	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	
$\Gamma(2K^+ \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{17}/\Gamma$
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.26</b>	90	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV

$\Gamma(2K^- \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.06	90	DOBBS	09	CLEO $e^+e^-$ at 4170 MeV	

### Leptonic and semileptonic modes

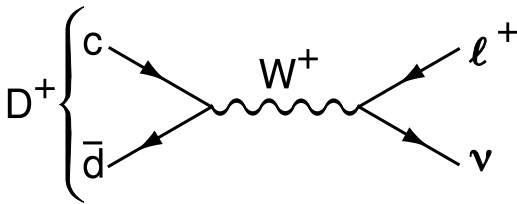
## LEPTONIC DECAYS OF CHARGED PSEUDO-SCALAR MESONS

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We review the physics of purely leptonic decays of  $\pi^\pm$ ,  $K^\pm$ ,  $D^\pm$ ,  $D_s^\pm$ , and  $B^\pm$  pseudoscalar mesons. The measured decay rates are related to the product of the relevant weak-interaction-based CKM matrix element of the constituent quarks and a strong interaction parameter related to the overlap of the quark and antiquark wave-functions in the meson, called the decay constant  $f_P$ . The leptonic decay constants for  $\pi^\pm$ ,  $K^\pm$ ,  $D^\pm$ ,  $D_s^\pm$ , and  $B^\pm$  mesons can be obtained with controlled theoretical uncertainties and high precision from *ab initio* lattice-QCD simulations. The combination of experimental leptonic decay-rate measurements and theoretical decay-constant calculations enables the determination of several elements of the CKM matrix within the standard model. These determinations are competitive with those obtained from semileptonic decays, and also complementary because they are sensitive to axial-vector (as opposed to vector) quark flavor-changing currents. They can also be used to test the unitarity of the first and second rows of the CKM matrix. Conversely, taking the CKM elements predicted by unitarity, one can infer “experimental” values for  $f_P$  that can be compared with theory. These provide tests of lattice-QCD methods, provided new-physics contributions to leptonic decays are negligible at the current level of precision. This review was prepared for the Particle Data Group’s 2016 edition, updating the versions in Refs. 1–3.

## I. INTRODUCTION

Charged mesons formed from a quark and an antiquark can decay to a charged lepton pair when these objects annihilate via a virtual  $W$  boson. Fig. 1 illustrates this process for the purely leptonic decay of a  $D^+$  meson.



**Figure 1:** The annihilation process for pure  $D^+$  leptonic decays in the Standard Model.

Similar quark-antiquark annihilations via a virtual  $W^+$  to the  $\ell^+\nu$  final states occur for the  $\pi^+$ ,  $K^+$ ,  $D_s^+$ , and  $B^+$  mesons. (Whenever pseudoscalar-meson charges are specified in this article, use of the charge-conjugate particles and corresponding decays are also implied.) Let  $P$  be any of these pseudoscalar mesons. To lowest order, the decay width is

$$\Gamma(P \rightarrow \ell\nu) = \frac{G_F^2}{8\pi} f_P^2 m_\ell^2 M_P \left(1 - \frac{m_\ell^2}{M_P^2}\right)^2 |V_{q_1 q_2}|^2 \quad (1)$$

Here  $M_P$  is the  $P$  mass,  $m_\ell$  is the  $\ell$  mass,  $V_{q_1 q_2}$  is the Cabibbo-Kobayashi-Maskawa (CKM) matrix element between the constituent quarks  $q_1 \bar{q}_2$  in  $P$ , and  $G_F$  is the Fermi coupling constant. The decay constant  $f_P$  is proportional to the matrix element of the axial current between the one- $P$ -meson state and the vacuum:

$$\langle 0 | \bar{q}_1 \gamma_\mu \gamma_5 q_2 | P(p) \rangle = i p_\mu f_P, \quad (2)$$

and can be thought of as the “wavefunction overlap” of the quark and antiquark. In this article we use the convention in which  $f_\pi \approx 130$  MeV.

The decay  $P^\pm$  starts with a spin-0 meson, and ends up with a left-handed neutrino or right-handed antineutrino. By angular momentum conservation, the  $\ell^\pm$  must then also be left-handed or right-handed, respectively. In the  $m_\ell = 0$  limit, the decay is forbidden, and can only occur as a result of the finite  $\ell$  mass. This helicity suppression is the origin of the  $m_\ell^2$  dependence of the decay width. Radiative corrections are needed when the final charged particle is an electron or muon; for the  $\tau$  they are greatly suppressed due to the large lepton mass, and hence negligible.

Measurements of purely leptonic decay branching fractions and lifetimes allow an experimental determination of the product  $|V_{q_1 q_2}| f_P$ . If the decay constant  $f_P$  is known to sufficient precision from theory, one can obtain the corresponding CKM element within the standard model. If, on the other hand, one takes the value of  $|V_{q_1 q_2}|$  assuming CKM unitarity, one can infer an “experimental measurement” of the decay constant that can then be compared with theory.

The importance of measuring  $\Gamma(P \rightarrow \ell\nu)$  depends on the particle being considered. Leptonic decays of charged pseudoscalar mesons occur at tree level within the standard model. Thus one does not expect large new-physics contributions to measurements of  $\Gamma(P \rightarrow \ell\nu)$  for the lighter mesons  $P = \pi^+, K^+$ , and these processes in principle provide clean standard-model determinations of  $V_{ud}$  and  $V_{us}$ . The situation is different for leptonic decays of charm and bottom mesons. The presence of new heavy particles such as charged Higgs bosons or leptoquarks could lead to observable effects in  $\Gamma(P \rightarrow \ell\nu)$  for  $P = D_{(s)}^+, B^+$  [4–8]. Thus the determination of  $|V_{ub}|$  from  $B^+ \rightarrow \tau\nu$  decay, in particular, should be considered a probe of new physics. More generally, the ratio of leptonic decays to  $\tau\nu$  over  $\mu\nu$  final states probes lepton universality [4,9].

The determinations of CKM elements from leptonic decays of charged pseudoscalar mesons provide complementary

# Meson Particle Listings

$D_s^\pm$

information to those from other decay processes. The decay  $P \rightarrow \ell\nu$  proceeds in the standard model via the axial-vector current  $\bar{q}_1\gamma_\mu\gamma_5q_2$ , whereas semileptonic pseudoscalar meson decays  $P_1 \rightarrow P_2\ell\nu$  proceed via the vector current  $\bar{q}_1\gamma_\mu q_2$ . Thus the comparison of determinations of  $|V_{q_1q_2}|$  from leptonic and semileptonic decays tests the  $V - A$  structure of the standard-model electroweak charged-current interaction. More generally, a small right-handed admixture to the standard-model weak current would lead to discrepancies between  $|V_{q_1q_2}|$  obtained from leptonic pseudoscalar-meson decays, exclusive semileptonic pseudoscalar-meson decays, exclusive semileptonic baryon decays, and inclusive semileptonic decays [10,11].

Both measurements of the decay rates  $\Gamma(P \rightarrow \ell\nu)$  and theoretical calculations of the decay constants  $f_P$  for  $P = \pi^+, K^+, D_{(s)}^+$  from numerical lattice-QCD simulations are now quite precise. As a result, the elements of the first row of the CKM matrix  $|V_{ud}|$  and  $|V_{us}|$  can be obtained to sub-percent precision from  $\pi^+ \rightarrow \ell\nu$  and  $K^+ \rightarrow \ell\nu$ , where the limiting error is from theory. The elements of the second row of the CKM matrix  $|V_{cd(s)}|$  can be obtained from leptonic decays of charged pseudoscalar mesons to few-percent precision, where here the limiting error is from experiment. These enable stringent tests of the unitarity of the first and second rows of the CKM matrix.

This review is organized as follows. Because the experimental and theoretical issues associated with measurements of pions and kaons, charmed mesons, and bottom mesons differ, we discuss each one separately. We begin with the pion and kaon system in Sec. II. First, in Sec. II.A we review current measurements of the experimental decay rates. We provide tables of branching-ratio measurements and determinations of the product  $|V_{ud(s)}|f_{\pi^+(K^+)}$ , as well as average values for these quantities including correlations and other effects needed to combine results. Then, in Sec. II.B we summarize the status of theoretical calculations of the decay constants. We provide tables of recent lattice-QCD results for  $f_{\pi^+}$ ,  $f_{K^+}$ , and their ratio from simulations including dynamical  $u, d, s$ , and (in some cases  $c$ ) quarks, and present averages for each of these quantities including correlations and strong SU(2)-isospin corrections as needed. We note that, for the leptonic decay constants in Sec. II.B, Sec. III.B, and Sec. IV.B, when available we use preliminary averages from the Flavor Lattice Averaging Group [12,13] that update the determinations in Ref. 14 to include results that have appeared since their most recent review, which dates from 2013. We next discuss the charmed meson system in Sec. III, again reviewing current experimental rate measurements in Sec. III.A and theoretical decay-constant calculations in Sec. III.B. Last, we discuss the bottom meson system in Sec. IV, following the same organization as the two previous sections.

After having established the status of both experimental measurements and theoretical calculations of leptonic charged pseudoscalar-meson decays, we discuss some implications for phenomenology in Sec. V. We combine the average  $\mathcal{B}(P \rightarrow \ell\nu)$  with the average  $f_P$  to obtain the relevant CKM elements from

leptonic decays, and then compare them with determinations from other processes. We also use the CKM elements obtained from leptonic decays to test the unitarity of the first and second rows of the CKM matrix. Further, as in previous reviews, we combine the experimental  $\mathcal{B}(P \rightarrow \ell\nu)$ s with the associated CKM elements obtained from CKM unitarity to infer “experimental” values for the decay constants; the comparison with theory provides a test of lattice and other QCD approaches assuming that new-physics contributions to these processes are not significant.

## II. PIONS AND KAONS

### A. Experimental rate measurements

The leading-order expression for  $\Gamma(P \rightarrow \ell\nu)$  in Eq. (1) is modified by radiative corrections arising from diagrams involving photons, in some cases with additional quark loops. These electroweak and “hadronic” contributions can be combined into an overall factor that multiplies the rate in the presence of only the strong interaction ( $\Gamma^{(0)}$ ) as follows (cf. Refs. 15,16, and references therein):

$$\Gamma(P \rightarrow \ell\nu) = \Gamma^{(0)} \left[ 1 + \frac{\alpha}{\pi} C_P \right], \quad (3)$$

where  $C_P$  differs for  $P = \pi, K$ . The inclusion of these corrections is numerically important given the level of precision achieved on the experimental measurements of the  $\pi^\pm \rightarrow \mu^\pm\nu$  and  $K^\pm \rightarrow \mu^\pm\nu$  decay widths. The explicit expression for the term in brackets above including all known electroweak and hadronic contributions is given in Eq. (114) of Ref. 17. It includes the universal short-distance electroweak correction obtained by Sirlin [18], the universal long-distance correction for a point-like meson from Kinoshita [19], and corrections that depend on the hadronic structure [20]. We evaluate  $\delta_P \equiv (\alpha/\pi)C_P$  using the latest experimentally-measured meson and lepton masses and coupling constants from the Particle Data Group [3], and taking the low-energy constants (LECs) that parameterize the hadronic contributions from Refs. 17,21,22. The finite non-logarithmic parts of the LECs were estimated within the large- $N_C$  approximation assuming that contributions from the lowest-lying resonances dominate. We therefore conservatively assign a 100% uncertainty to the LECs, which leads to a  $\pm 0.9$  error in  $C_{\pi,K}$ .<sup>1</sup> We obtain the following correction factors to the individual charged pion and kaon decay widths:

$$\delta_\pi = 0.0176(21) \quad \text{and} \quad \delta_K = 0.0107(21). \quad (4)$$

<sup>1</sup> This uncertainty on  $C_{\pi,K}$  is smaller than the error estimated by Marciano and Sirlin in Ref. 23, which predates the calculations of the hadronic-structure contributions in Refs. 17, 20–22. The hadronic LECs incorporate the large short-distance electroweak logarithm discussed in Ref. 23, and their dependence on the chiral renormalization scale cancels the scale-dependence induced by chiral loops, thereby removing the dominant scale uncertainty of the Marciano–Sirlin analysis [23].

The error on the ratio of kaon-to-pion leptonic decay widths is under better theoretical control because the hadronic contributions from low-energy constants estimated within the large- $N_c$  framework cancel at lowest order in the chiral expansion. For the ratio, we use the correction factor

$$\delta_{K/\pi} = -0.0069(17), \quad (5)$$

where we take the estimated error due to higher-order corrections in the chiral expansion from Ref. 24.

The sum of branching fractions for  $\pi^- \rightarrow \mu^- \bar{\nu}$  and  $\pi^- \rightarrow \mu^- \bar{\nu} \gamma$  is 99.98770(4)% [3]. The two modes are difficult to separate experimentally, so we use this sum. Together with the lifetime 26.033(5) ns [3] this implies  $\Gamma(\pi^- \rightarrow \mu^- \bar{\nu}[\gamma]) = 3.8408(7) \times 10^7 \text{ s}^{-1}$ . The right-hand side of Eq. (1) is modified by the factor  $1.0176 \pm 0.0021$  mentioned above to include photon emission and radiative corrections [23,25]. The decay rate together with the masses from the 2014 PDG review [3] gives

$$f_{\pi^-} |V_{ud}| = (127.13 \pm 0.02 \pm 0.13) \text{ MeV}, \quad (6)$$

where the errors are from the experimental rate measurement and the radiative correction factor  $\delta_\pi$  in Eq. (4), respectively. The uncertainty is dominated by that from theoretical estimate of the hadronic structure-dependent radiative corrections, which include next-to-leading order contributions of  $\mathcal{O}(e^2 p_{\pi,K}^2)$  in chiral perturbation theory [17].

The data on  $K_{\mu 2}$  decays have been updated recently through a global fit to branching ratios and lifetime measurements [26]:  $\mathcal{B}(K^- \rightarrow \mu^- \bar{\nu}[\gamma]) = 63.58(11)\%$  and  $\tau_{K^\pm} = 12.384(15) \text{ ns}$ . The improvement in the branching ratio is primarily due to a new measurement of  $\mathcal{B}(K^\pm \rightarrow \pi^\pm \pi^+ \pi^-)$  from KLOE-2 [27], which is correlated with  $\mathcal{B}(K_{\mu 2}^\pm)$  through the constraint that the sum of individual branching ratios must equal unity. The sum of branching fractions for  $K^- \rightarrow \mu^- \bar{\nu}$  and  $K^- \rightarrow \mu^- \bar{\nu} \gamma$  and the lifetime imply  $\Gamma(K^- \rightarrow \mu^- \bar{\nu}[\gamma]) = 5.134(11) \times 10^7 \text{ s}^{-1}$ . Again taking the 2014 PDG masses [3], this decay rate implies

$$f_{K^+} |V_{us}| = (35.09 \pm 0.04 \pm 0.04) \text{ MeV}, \quad (7)$$

where the errors are from the experimental rate measurement and the radiative correction factor  $\delta_K$ , respectively.

Short-distance radiative corrections cancel in the ratio of pion-to-kaon decay rates [28]:

$$\frac{\Gamma_{K_{\ell 2}[\gamma]}}{\Gamma_{\pi_{\ell 2}[\gamma]}} = \frac{|V_{us}|^2 f_{K^-}^2}{|V_{ud}|^2 f_{\pi^-}^2} \frac{m_K(1 - m_\ell^2/m_K^2)^2}{m_\pi(1 - m_\ell^2/m_\pi^2)^2} (1 + \delta_{K/\pi}), \quad (8)$$

where  $\delta_{K/\pi}$  is given in Eq. (5). The left-hand side of Eq. (8) is 1.3367(28), yielding

$$\frac{|V_{us}| f_{K^-}}{|V_{ud}| f_{\pi^-}} = 0.27599 \pm 0.00029 \pm 0.00024, \quad (9)$$

where the first uncertainty is due to the branching fractions and the second is due to  $\delta_{K/\pi}$ . Here the estimated error on the hadronic structure-dependent radiative corrections is commensurate with the experimental error.

In summary, the main experimental results pertaining to charged pion and kaon leptonic decays are

$$|V_{ud}| f_{\pi^-} = (127.13 \pm 0.02 \pm 0.13) \text{ MeV}, \quad (10)$$

$$|V_{us}| f_{K^+} = (35.09 \pm 0.04 \pm 0.04) \text{ MeV}, \quad (11)$$

$$\frac{|V_{us}| f_{K^+}}{|V_{ud}| f_{\pi^-}} = 0.27599 \pm 0.00029 \pm 0.00024, \quad (12)$$

where the errors are from the experimental uncertainties in the branching fractions and the theoretical uncertainties in the radiative correction factors  $\delta_P$ , respectively.

### B. Theoretical decay-constant calculations

Table 1 presents recent lattice-QCD calculations of the charged pion and kaon decay constants and their ratio from simulations with three ( $N_f = 2+1$ ) or four flavors ( $N_f = 2+1+1$ ) of dynamical quarks. The results have been obtained using several independent sets of gauge-field configurations, and a variety of lattice fermion actions that are sensitive to different systematic uncertainties.<sup>2</sup> The lattice-QCD uncertainties on both the individual decay constants and their ratio have now reached sub-percent precision. The SU(3)-breaking ratio  $f_{K^+}/f_{\pi^+}$  can be obtained with especially small errors because statistical errors associated with the Monte Carlo simulations are correlated between the numerator and denominator, as are some systematics. The good agreement between these largely independent determinations indicates that the lattice-QCD uncertainties are controlled and that the associated error estimates are reliable.<sup>3</sup>

Table 1 also shows the 2015 preliminary three- and four-flavor averages for the pion and kaon decay constants and their ratio from the Flavour Lattice Averaging Group (FLAG) [12,13] in the lines labeled “FLAG 15 average.” These preliminary updates of the 2013 FLAG averages [14] include only those results from Table 1 that are published in refereed journals, or that are straightforward conference updates of published analyses. In the (2+1+1)-flavor averages, the statistical errors of HPQCD and Fermilab/MILC were conservatively treated as 100% correlated because the calculations employed some of the same gauge-field configurations. The errors have also been increased by the  $\sqrt{\chi^2/\text{dof}}$  to reflect a slight tension between the results. There are no four-flavor lattice-QCD results for the pion decay constant in Table 1 because all of the calculations listed use the quantity  $f_{\pi^+}$  to fix the absolute lattice scale needed to convert from lattice-spacing units to GeV [31–33].

<sup>2</sup> See the PDG mini-review on “Lattice Quantum Chromodynamics” [29] for a general review of numerical lattice-QCD simulations. Details on the different methods used in modern lattice-QCD calculations are provided in Appendix A of the FLAG “Review of lattice results concerning low energy particle physics” [14].

<sup>3</sup> The recent review [30] summarizes the large body of evidence validating the methods employed in modern lattice-QCD simulations.

## Meson Particle Listings

 $D_s^\pm$ 

**Table 1:** Recent lattice-QCD results for  $f_{\pi^+}$ ,  $f_{K^+}$ , and their ratio. The upper and lower panels show  $(2+1+1)$ -flavor and  $(2+1)$ -flavor determinations, respectively. When two errors are shown, they are statistical and systematic, respectively. Results for  $f_\pi$  and  $f_K$  in the isospin-symmetric limit  $m_u = m_d$  are noted with an “\*”; they are corrected for isospin breaking via Eq. (13)–Eq. (15) before computing the averages. Unpublished results noted with a “†” or “‡” are not included in the averages.

Reference	$N_f$	$f_{\pi^+}(\text{MeV})$	$f_{K^+}(\text{MeV})$	$f_{K^+}/f_{\pi^+}$
ETM 14 [31] §	2+1+1	–	154.4(1.5)(1.3)	1.184(12)(11)
Fermilab/MILC 14 [32] §	2+1+1	–	155.92(13)( $^{+42}_{-34}$ )	1.1956(10)( $^{+26}_{-18}$ )
HPQCD 13 [33] §	2+1+1	–	155.37(20)(28)	1.1916(15)(16)
FLAG 15 average [12,13] ¶	2+1+1	–	155.6(0.4)	1.193(3)
RBC/UKQCD 14 [34] *,†	2+1	130.19(89)	155.51(83)	1.1945(45)
RBC/UKQCD 12 [35] *	2+1	127(3)(3)	152(3)(2)	1.199(12)(14)
Laiho & Van de Water 11 [36] ‡	2+1	130.53(87)(210)	156.8(1.0)(1.7)	1.202(11)(9)(2)(5)
MILC 10 [37]	2+1	129.2(0.4)(1.4)	156.1(4)( $^{+6}_{-9}$ )	1.197(2)( $^{+3}_{-7}$ )
BMW 10 [38] *	2+1	–	–	1.192(7)(6)
HPQCD/UKQCD 07 [39] *	2+1	132(2)	157(2)	1.189(2)(7)
FLAG 15 average [12,13] ¶	2+1	130.2(1.4)	155.9(0.9)	1.192(5)
Our average	Both	130.2(1.7)	155.6(0.4)	1.1928(26)

§ PDG 2014 value of  $f_{\pi^+} = 130.41(21)$  MeV used to set absolute lattice scale.

¶ Preliminary numbers shown here may change if further new lattice-QCD calculations are published before the deadline for inclusion in the final 2015 FLAG review.

† Preprint submitted to Phys. Rev. D. Published RBC/UKQCD 12 results included in  $N_f = 2+1$  average.

‡ Lattice 2011 conference proceedings.

All of the results in Table 1 were obtained using isospin-symmetric gauge-field configurations, *i.e.*, the dynamical up and down quarks have the same mass. Most calculations of pion and kaon decay constants now include the dominant effect of nondegenerate up- and down-quark masses by evaluating the masses of the constituent light (valence) quarks in the pion at the physical up- and down-quark masses, respectively, and evaluating the mass of the valence light quark in the kaon at the physical  $m_u$ . Those results obtained with degenerate up and down valence quarks are corrected for isospin breaking using chiral perturbation theory ( $\chi$ PT) before being averaged. The isospin-breaking corrections at next-to-leading order in  $\chi$ PT can be parameterized as [24,40]

$$f_\pi = f_{\pi^+}, \quad (13)$$

$$f_K = f_{K^+} (1 - \delta_{\text{SU}(2)}/2), \quad (14)$$

$$\frac{f_K}{f_\pi} = \frac{1}{\sqrt{\delta_{\text{SU}(2)} + 1}} \frac{f_{K^+}}{f_{\pi^+}} \quad (15)$$

where the expression for  $\delta_{\text{SU}(2)}$  in terms of the quark masses, meson masses, and decay constants, is given in Eq. (37) of Ref. 14. Numerically, values of  $\delta_{\text{SU}(2)} \approx -0.004$  were employed by FLAG to obtain the  $(2+1)$ -flavor averages in Table 1,

but some direct lattice-QCD calculations of  $\delta_{\text{SU}(2)}$  give larger values [31,33,41] and further studies are needed.

To obtain the best decay-constant values for comparison with experimental rate measurements and other phenomenological applications, we combine the available  $(2+1)$ - and  $(2+1+1)$ -flavor lattice-QCD results, first accounting for the omission of charm sea quarks in the three-flavor simulations. The error introduced by omitting charm sea quarks can be roughly estimated by expanding the charm-quark determinant in powers of  $1/m_c$  [42]; the resulting leading contribution is of order  $\alpha_s (\Lambda_{\text{QCD}}/2m_c)^2$  [43]. Taking the  $\overline{\text{MS}}$  values  $\overline{m}_c(\overline{m}_c) = 1.275$  GeV,  $\overline{\Lambda}_{\text{QCD}} \sim 340$  MeV from FLAG [14], and  $\overline{\alpha}(\overline{m}_c) \sim 0.4$ , leads to an estimate of about 0.7% for the contribution to the decay constants from charm sea quarks. The charm sea-quark contribution to ratios of decay constants is expected to be further suppressed by the  $\text{SU}(3)$ -breaking factor  $(m_s - m_d)/\Lambda_{\text{QCD}}$ , and hence about 0.2%.

We can compare these power-counting estimates of charm sea-quark contributions to the observed differences between the  $(2+1)$ - and  $(2+1+1)$ -flavor lattice-QCD averages for kaon,  $D_{(s)}$ -meson, and  $B_{(s)}$ -decay constants and ratios in Table 1, Table 4, and Table 6. Of these, the kaon decay constants have been calculated most precisely, and the two- and three- and four-flavor averages for  $f_{K^+}$  and  $f_{K^+}/f_{\pi^+}$  agree within sub-percent errors. Within present uncertainties, however, effects of this



size in pseudoscalar-meson decay constants cannot be ruled out. Therefore, to be conservative, in this review we add in quadrature additional systematic errors of 0.7% and 0.2% to all (2+1)-flavor decay-constant and decay-constant-ratio averages, respectively, to account for the omission of charm sea quarks. Numerically, this increases the errors by at most about 50% for  $f_{K^+}$  and less for all other decay constants and ratios, indicating that the published (2+1)-flavor lattice-QCD results and uncertainties are reliable.

Our final preferred theoretical values for the charged pion and kaon decay constants are

$$\begin{aligned} \text{Our averages : } f_{\pi^+} &= 130.2(1.7) \text{ MeV}, \\ f_{K^+} &= 155.6(0.4) \text{ MeV}, \\ \frac{f_{K^+}}{f_{\pi^+}} &= 1.1928(26), \end{aligned} \quad (16)$$

where  $f_{\pi^+}$  is simply the (2+1)-flavor FLAG average with the error increased by the estimated 0.7% charm sea-quark contribution. For  $f_{K^+}$  and  $f_{K^+}/f_{\pi^+}$ , we take a simple weighted average of the (2+1)- and (2+1+1)-flavor FLAG values, because they are each obtained from a sufficient number of independent calculations that we do not expect there to be significant correlations. In practice, the addition of the charm sea-quark error has a tiny impact on our final values in Eq. (16), increasing the uncertainty on  $f_{\pi^+}$  by 0.3 MeV, and the central value for  $f_{K^+}/f_{\pi^+}$  by one in the last digit.

### III. CHARMED MESONS

#### A. Experimental rate measurements

Measurements have been made for  $D^+ \rightarrow \mu^+\nu$ ,  $D_s^+ \rightarrow \mu^+\nu$ , and  $D_s^+ \rightarrow \tau^+\nu$ . Only an upper limit has been determined for  $D^+ \rightarrow \tau^+\nu$ . Both CLEO-c and BES have made measurements of  $D^+$  decay using  $e^+e^-$  collisions at the  $\psi(3770)$  resonant energy where  $D^-D^+$  pairs are copiously produced. They fully reconstruct one of the  $D$ 's, say the  $D^-$ . Counting the number of these events provides the normalization for the branching fraction measurement. They then find a candidate  $\mu^+$ , and then form the missing-mass squared,  $MM^2 = (E_{\text{CM}} - E_{D^-})^2 - (\vec{p}_{\text{CM}} - \vec{p}_{D^-} - \vec{p}_{\mu^+})^2$ , taking into account their knowledge of the center-of-mass energy,  $E_{\text{CM}}$ , and momentum,  $p_{\text{CM}}$ , that equals zero in  $e^+e^-$  collisions. A peak at zero  $MM^2$  implies the existence of a missing neutrino and hence the  $\mu^+\nu$  decay of the  $D^+$ . CLEO-c does not explicitly identify the muon, so their data consists of a combination of  $\mu^+\nu$  and  $\tau^+\nu$ ,  $\tau^+ \rightarrow \pi^+\nu$  events. This permits them to do two fits: in one they fit for the individual components, and in the other they fix the ratio of  $\tau^+\nu/\mu^+\nu$  events to be that given by the standard-model expectation. Thus, the latter measurement should be used for standard-model comparisons and the other for new-physics searches. Our average uses the fixed ratio value. The measurements are shown in Table 2.

**Table 2:** Experimental results for  $\mathcal{B}(D^+ \rightarrow \mu^+\nu)$ ,  $\mathcal{B}(D^+ \rightarrow \tau^+\nu)$ , and  $|V_{cd}|f_{D^+}$ . Numbers for  $|V_{cd}|f_{D^+}$  have been extracted using updated values for masses (see text). Radiative corrections are included. Systematic uncertainties arising from the  $D^+$  lifetime and mass are included. For the average  $\mu^+\nu$  number we use the CLEO-c result for  $\mu^+\nu + \tau^+\nu$ .

Experiment	Mode	$\mathcal{B}$	$ V_{cd} f_{D^+}$ (MeV)
CLEO-c [44,45]	$\mu^+\nu$	$(3.93 \pm 0.35 \pm 0.09) \times 10^{-4}$	$47.07 \pm 2.10 \pm 0.57$
CLEO-c [44,45]	$\mu^+\nu + \tau^+\nu$	$(3.82 \pm 0.32 \pm 0.09) \times 10^{-4}$	$46.41 \pm 1.94 \pm 0.57$
BES [46]	$\mu^+\nu$	$(3.71 \pm 0.19 \pm 0.06) \times 10^{-4}$	$45.73 \pm 1.17 \pm 0.38$
Our average	Lines 2+3	$(3.74 \pm 0.17) \times 10^{-4}$	$45.91 \pm 1.05$
CLEO-c [47,48]	$\tau^+\nu$	$< 1.2 \times 10^{-3}$	

To extract the value of  $|V_{cd}|f_{D^+}$  we use the well-measured  $D^+$  lifetime of 1.040(7) ps. The  $\mu^+\nu$  results include a 1% correction (lowering) of the rate due to the presence of the radiative  $\mu^+\nu\gamma$  final state based on the estimate by Dobrescu and Kronfeld [8].

We now discuss the  $D_s^+$ . Measurements of the leptonic decay rate have been made by several groups and are listed in Table 3 [47–53]. We exclude older values obtained by normalizing to  $D_s^+$  decay modes that are not well defined. Many measurements, for example, used the  $\phi\pi^+$  mode. This decay is a subset of the  $D_s^+ \rightarrow K^+K^-\pi^+$  channel which has interferences from other modes populating the  $K^+K^-$  mass region near the  $\phi$ , the most prominent of which is the  $f_0(980)$ . Thus the extraction of the effective  $\phi\pi^+$  rate is sensitive to the mass resolution of the experiment and the cuts used to define the  $\phi$  mass region [54].<sup>4</sup>

To find decays in the  $\mu^+\nu$  signal channels, CLEO, BaBar and Belle rely on fully reconstructing all the final state particles except for neutrinos and using a missing-mass technique to infer the existence of the neutrino. CLEO uses  $e^+e^- \rightarrow D_s D_s^*$  collisions at 4170 MeV, while Babar and Belle use  $e^+e^- \rightarrow DK n \pi D_s^*$  collisions at energies near the  $\Upsilon(4S)$ . CLEO does a similar analysis as was done for the  $D^+$  above. Babar and Belle do a similar  $MM^2$  calculation by using the reconstructed hadrons, the photon from the  $D_s^{*+}$  decay and a detected  $\mu^+$ . To get the normalization they do a  $MM^2$  fit without the  $\mu^+$  and use the signal at the  $D_s^+$  mass squared to determine the total  $D_s^+$  yield.

When selecting the  $\tau^+ \rightarrow \pi^+\bar{\nu}$  and  $\tau^+ \rightarrow \rho^+\bar{\nu}$  decay modes, CLEO uses both the calculation of the missing mass and the fact that there should be no extra energy in the event beyond that deposited by the measured tagged  $D_s^-$  and the  $\tau^+$  decay products. The  $\tau^+ \rightarrow e^+\nu\bar{\nu}$  mode, however, uses only extra energy. Babar and Belle also use the extra energy to discriminate signal from background in their  $\tau^+\nu$  measurements.

<sup>4</sup> We have not included the BaBar result for  $\mathcal{B}(D_s^+ \rightarrow \mu^+\nu)$  reported in Ref. 55 because this measurement determined the ratio of the leptonic decay rate to the hadronic decay rate  $\Gamma(D_s^+ \rightarrow \ell^+\nu)/\Gamma(D_s^+ \rightarrow \phi\pi^+)$ .

## Meson Particle Listings

 $D_s^\pm$ 

**Table 3:** Experimental results for  $\mathcal{B}(D_s^+ \rightarrow \mu^+\nu)$ ,  $\mathcal{B}(D_s^+ \rightarrow \tau^+\nu)$ , and  $|V_{cs}|f_{D_s^+}$ . Numbers for  $|V_{cs}|f_{D_s^+}$  have been extracted using updated values for masses (see text). The systematic uncertainty for correlated error on the  $D_s^+$  lifetime is included. The mass uncertainties are also common, but negligible. Common systematic errors in each experiment have been taken into account in the averages.

Experiment	Mode	$\mathcal{B}(\%)$	$ V_{cs} f_{D_s^+}$ (MeV)
CLEO-c [47,48]	$\mu^+\nu$	$0.565 \pm 0.045 \pm 0.017$	$250.8 \pm 10.0 \pm 4.2$
BaBar <sup>a</sup> [53]	$\mu^+\nu$	$0.602 \pm 0.038 \pm 0.034$	$258.9 \pm 8.2 \pm 7.5$
Belle [49]	$\mu^+\nu$	$0.531 \pm 0.028 \pm 0.020$	$243.1 \pm 6.4 \pm 4.9$
Our average	$\mu^+\nu$	$0.556 \pm 0.024$	$248.8 \pm 5.8$
CLEO-c [47,48]	$\tau^+\nu$ ( $\pi^+\bar{\nu}$ )	$6.42 \pm 0.81 \pm 0.18$	$270.8 \pm 17.1 \pm 4.2$
CLEO-c [50]	$\tau^+\nu$ ( $\rho^+\bar{\nu}$ )	$5.52 \pm 0.57 \pm 0.21$	$251.1 \pm 13.0 \pm 5.1$
CLEO-c [51,52]	$\tau^+\nu$ ( $e^+\nu\bar{\nu}$ )	$5.30 \pm 0.47 \pm 0.22$	$246.1 \pm 10.9 \pm 5.4$
BaBar [53]	$\tau^+\nu$ ( $e^+(\mu^+)\nu\bar{\nu}$ )	$5.00 \pm 0.35 \pm 0.49$	$239.0 \pm 8.4 \pm 11.9$
Belle [49]	$\tau^+\nu$ ( $\pi^+\bar{\nu}$ )	$6.04 \pm 0.43^{+0.46}_{-0.40}$	$262.7 \pm 9.3^{+10.2}_{-8.9}$
Belle [49]	$\tau^+\nu$ ( $e^+\nu\bar{\nu}$ )	$5.37 \pm 0.33^{+0.35}_{-0.31}$	$247.7 \pm 7.6^{+8.3}_{-7.4}$
Belle [49]	$\tau^+\nu$ ( $\mu^+\nu\bar{\nu}$ )	$5.86 \pm 0.37^{+0.34}_{-0.59}$	$258.7 \pm 8.2^{+7.7}_{-13.2}$
Our average	$\tau^+\nu$	$5.56 \pm 0.22$	$252.1 \pm 5.2$
Our average	$\mu^+\nu + \tau^+\nu$		$250.9 \pm 4.0$

<sup>a</sup>We do not use a previous unpublished BaBar result from a subsample of data that uses a different technique for obtaining the branching fraction normalization [56].

We extract the decay constant times the CKM factor from the measured branching ratios using the  $D_s^+$  mass of 1.96830(11) GeV, the  $\tau^+$  mass of 1.77682(16) GeV, and a  $D_s^+$  lifetime of 0.500(7) ps [3]. CLEO has included the radiative correction of 1% in the  $\mu^+\nu$  rate listed in the Table [8] (the  $\tau^+\nu$  rates need not be corrected). Other theoretical calculations show that the  $\gamma\mu^+\nu$  rate is a factor of 40–100 below the  $\mu^+\nu$  rate for charm [57–66]. As this is a small effect we do not attempt to correct the other measurements. The values for  $f_{D_s^+}|V_{cs}|$  are in good agreement for the two decay modes. Our average value including both the  $\mu^+\nu$  and  $\tau^+\nu$  final states is  $250.9 \pm 4.0$  MeV.

### B. Theoretical decay-constant calculations

Table 4 presents recent theoretical calculations of the charged  $D^{*-}$  and  $D_s$ -meson decay constants and their ratio. The upper two panels show results from lattice-QCD simulations with three ( $N_f = 2 + 1$ ) or four flavors ( $N_f = 2 + 1 + 1$ ) of dynamical quarks. Although there are fewer available results than for the pion and kaon sector, both  $f_{D^+}$  and  $f_{D_s}$  have been obtained using multiple sets of gauge-field configurations with different lattice fermion actions, providing independent confirmation. For comparison, the bottom panel of Table 4 shows non-lattice determinations from QCD sum rules and the light-front quark model; only results which include uncertainty estimates are shown. The lattice and non-lattice results agree, but the uncertainties on  $D_{(s)}^+$ -meson decay constants from lattice QCD have now reached significantly greater precision than those from other approaches.

The lattice-QCD results in Table 4 were all obtained using isospin-symmetric gauge-field configurations. The two calculations by the Fermilab Lattice and MILC Collaborations [69,32], however, include the dominant strong isospin-breaking contribution by evaluating the mass of the valence light quark in the  $D^+$ -meson decay constant at the physical down-quark mass. Reference 32 provides a determination of the size of this correction,

$$f_{D^+} - f_D = 0.47(1)_{(-6)}^{(+25)} \text{ MeV}, \quad (17)$$

where  $f_D$  is the value of the  $D$ -meson decay constant evaluated at the average up-down quark mass. Eq. (17) implies that the correction to the  $SU(3)_f$ -breaking ratio is

$$\frac{f_{D_s}}{f_{D^+}} - \frac{f_{D_s}}{f_D} = -0.0026, \quad (18)$$

taking the central values for  $f_{D^+}$  and  $f_{D_s}$  from the same work. Because the errors on the calculations listed in Table 4 that neglect isospin breaking are still about  $5\text{--}8 \times$  larger than the sizes of the shifts in Eqs. (17)–(18), we do not correct any results *a posteriori* for this effect in the current review. Nevertheless, we strongly encourage future lattice-QCD publications to present results for both the  $D^{*-}$  and  $D^0$ -meson decay constants. Including the effect of isospin breaking will be essential once lattice-QCD calculations of  $f_D$  and  $f_{D_s}/f_D$  reach the level of precision in Eqs. (17)–(18).

We average the lattice-QCD results in Table 4 accounting for possible correlations between them following the approach established by Laiho *et al.* [77]. Whenever we have reason to believe that a source of uncertainty is correlated between two

**Table 4:** Recent theoretical determinations of  $f_{D^+}$ ,  $f_{D_s}$ , and their ratio. The upper panels show results from lattice-QCD simulations with  $(2+1+1)$  and  $(2+1)$  dynamical quark flavors, respectively. Statistical and systematic errors are quoted separately. Lattice-QCD results for  $f_D$  and  $f_{D_s}/f_D$  in the isospin-symmetric limit  $m_u = m_d$  are noted with an “\*”. The bottom panel shows estimates from QCD sum rules (QCD SR) and the light-front quark model (LFQM). These are not used to obtain our preferred decay-constant values.

Reference	Method	$N_f$	$f_{D^+}(\text{MeV})$	$f_{D_s}(\text{MeV})$	$f_{D_s}/f_{D^+}$
ETM 14 [31] *	LQCD	2+1+1	207.4(3.7)(0.9)	247.2(3.9)(1.4)	1.192(19)(11)
Fermilab/MILC 14 [32]	LQCD	2+1+1	212.6(0.4)( $^{+1.0}_{-1.2}$ )	249.0(0.3)( $^{+1.1}_{-1.5}$ )	1.1712(10)( $^{+29}_{-32}$ )
Average	LQCD	2+1+1	212.2(1.5)	248.8(1.3)	1.172(3)
$\chi$ QCD 14 [67] *	LQCD	2+1	—	254(2)(4)	—
HPQCD 12 [68] *	LQCD	2+1	208.3(1.0)(3.3)	—	1.187(4)(12)
Fermilab/MILC 11 [69]	LQCD	2+1	218.9(9.2)(6.6)	260.1(8.9)(6.1)	1.188(14)(21)
HPQCD 10 [70] *	LQCD	2+1	—	248.0(1.4)(2.1)	—
Average	LQCD	2+1	209.2(3.3)	249.8(2.3)	1.187(12)
Our average	LQCD	Both	211.9(1.1)	249.0(1.2)	1.173(3)
Wang 15 [71] §	QCD SR		208(10)	240(10)	1.15(6)
Gelhausen 13 [72]	QCD SR		201( $^{+12}_{-13}$ )	238( $^{+13}_{-23}$ )	1.15( $^{+0.04}_{-0.05}$ )
Narison 12 [73]	QCD SR		204(6)	246(6)	1.21(4)
Lucha 11 [74]	QCD SR		206.2(8.9)	245.3(16.3)	1.193(26)
Hwang 09 [75]	LFQM		—	264.5(17.5)¶	1.29(7)

§ Obtained using  $m_c^{\overline{\text{MS}}}$ ; results using  $m_c^{\text{pole}}$  are also given in the paper.

¶ Obtained by combining PDG value  $f_D = 205.8(8.9)$  MeV [76] with  $f_{D_s}/f_D$  from this work.

results, we conservatively take the correlation to be 100% when calculating the average. We then construct the correlation matrix for the set of lattice-QCD results using the prescription of Schmelling [78].

We first separately average the three- and four-flavor results for the charged  $D_{(s)}^+$ -meson decay constants and their ratio. There have been no new three-flavor lattice-QCD calculations of  $f_{D^+}$  or  $f_{D_s^+}/f_{D^+}$  since 2013, so we take the  $(2+1)$ -flavor averages from FLAG [14]. In this average, the statistical errors were treated as 100% correlated between the results of Fermilab/MILC [69] and HPQCD [68] because the calculations employed some of the same ensembles of gauge-field configurations. For  $f_{D_s}$ , we average the  $(2+1)$ -flavor results given in Table 4, again treating the Fermilab/MILC [69] and HPQCD [70] statistical errors as correlated, and taking the  $\chi$ QCD result [67] to be independent. For the  $(2+1+1)$ -flavor  $D_{(s)}$ -meson decay constants, we take a simple weighted average of the ETM [31] and Fermilab/MILC 14 results [32] in Table 4. We expect them to be independent because the calculations use different light-quark and gluon actions and different treatments of the chiral-continuum extrapolation. Our separate three- and four-flavor averages are listed in the lines labeled “Average” in Table 4, where the errors on the  $(2+1)$ -flavor  $f_{D_s}$  and

$(2+1+1)$ -flavor  $f_D$  averages have been rescaled by the factors  $\sqrt{(\chi^2/\text{dof})} = 1.1$  and  $\sqrt{(\chi^2/\text{dof})} = 1.3$ , respectively.<sup>5</sup>

To obtain the single-best values of the  $D_{(s)}^+$ -meson decay constants for phenomenology applications, we combine the available  $(2+1)$ - and  $(2+1+1)$ -flavor lattice-QCD results, which are compatible within the current level of precision. We account for the omission of charm sea-quark contributions in the three-flavor calculations by adding to the errors on the  $(2+1)$ -flavor averages in Table 4 our power-counting estimates of charm sea-quark errors from Sec. II.B. Because the estimated charm sea-quark errors of 0.7% for decay constants and 0.2% for decay-constant ratios are less than those on the  $(2+1)$ -flavor averages, adding them in quadrature has a small impact on the total uncertainties. The error increase is at most about 25% for  $f_{D_s}$ , and below 10% for both  $f_{D^+}$  and  $f_{D_s}/f_{D^+}$ . Our final preferred theoretical values for the charged  $D_{(s)}^+$ -meson decay constants are given by the weighted average of the entries in the two lines labeled “Average” in Table 4, after including the additional charm sea-quark errors in the  $(2+1)$ -flavor entries:

Our averages :  $f_{D^+} = 211.9(1.1)$  MeV ,

<sup>5</sup> After this article was submitted for review, preliminary  $(2+1)$ - and  $(2+1+1)$ -flavor FLAG averages for  $f_D$ ,  $f_{D_s}$ , and  $f_{D_s}/f_D$  were presented in Ref. 79 that are identical to our separate averages in Table 4.

# Meson Particle Listings

$D_s^\pm$

$$\begin{aligned} f_{D_s} &= 249.0(1.2) \text{ MeV}, \\ \frac{f_{D_s}}{f_{D^+}} &= 1.173(3). \end{aligned} \quad (19)$$

In practice, the errors on the (2+1+1)-flavor averages are so much smaller than on the (2+1)-flavor averages that the combination in Eq. (19) is almost identical to the (2+1+1)-flavor average in Table 4. The most precise result from Fermilab/MILC, in particular, has a large weight in the average.

## IV. BOTTOM MESONS

### A. Experimental rate measurements

The Belle and BaBar collaborations have found evidence for  $B^- \rightarrow \tau^- \bar{\nu}$  decay in  $e^+e^- \rightarrow B^- B^+$  collisions at the  $\Upsilon(4S)$  energy. The analysis relies on reconstructing a hadronic or semileptonic  $B$  decay tag, finding a  $\tau$  candidate in the remaining track and photon candidates, and examining the extra energy in the event which should be close to zero for a real  $\tau^-$  decay to  $e^- \nu \bar{\nu}$  or  $\mu^- \nu \bar{\nu}$  opposite a  $B^+$  tag. While the BaBar results have remained unchanged, Belle reanalyzed both samples of their data. The branching fraction using hadronic tags changed from  $1.79^{+0.56+0.46}_{-0.49-0.51} \times 10^{-4}$  [80] to  $0.72^{+0.27}_{-0.25} \pm 0.11 \times 10^{-4}$  [81], while the corresponding change using semileptonic tags was from  $1.54^{+0.38+0.29}_{-0.37-0.31}$  to  $1.25 \pm 0.28 \pm 0.27$ . These changes demonstrate the difficulty of the analysis. The results are listed in Table 5.

There are large backgrounds under the signals in all cases. The systematic errors are also quite large. Thus, the significances are not that large. Belle quotes  $4.6\sigma$  for their combined hadronic and semileptonic tags, while BaBar quotes  $3.3\sigma$  and  $2.3\sigma$ , for hadronic and semileptonic tags. Greater precision is necessary to determine if any effects beyond the Standard Model are present.

**Table 5:** Experimental results for  $\mathcal{B}(B^- \rightarrow \tau^- \bar{\nu})$  and  $|V_{ub}|f_{B^+}$ .

Experiment	Tag	$\mathcal{B}$ (units of $10^{-4}$ )	$ V_{ub} f_{B^+}$ (MeV)
Belle [81]	Hadronic	$0.72^{+0.27}_{-0.25} \pm 0.11$	
Belle [82]	Semileptonic	$1.25 \pm 0.28 \pm 0.27$	
Belle [82]	Average	$0.91 \pm 0.22$	$0.72 \pm 0.09$
BaBar [83]	Hadronic	$1.83^{+0.53}_{-0.49} \pm 0.24$	
BaBar [84]	Semileptonic	$1.7 \pm 0.8 \pm 0.2$	
BaBar [83]	Average	$1.79 \pm 0.48$	$1.01 \pm 0.14$
Our average		$1.06 \pm 0.20$	$0.77 \pm 0.07$

To extract the value of  $|V_{ub}|f_{B^+}$  we use the PDG 2014 value of the  $B^+$  lifetime of  $1.638 \pm 0.004$  ps, and the  $\tau^+$  and  $B^+$  masses of 1.77684 and 5.27926 GeV, respectively.

### B. Theoretical decay-constant calculations

Table 6 and Table 7 present theoretical calculations of the  $B^+$ -,  $B^0$ -, and  $B_s$ -meson decay constants and their ratios. (The decay constants of the neutral  $B^0$  and  $B_s$  mesons enter the rates for the rare leptonic decays  $B_{d,s} \rightarrow \mu^+ \mu^-$ .) The upper

two panels show results from lattice-QCD simulations with three ( $N_f = 2+1$ ) or four flavors ( $N_f = 2+1+1$ ) of dynamical quarks. For all decay constants, calculations using different gauge-field configurations, light-quark actions, and  $b$ -quark actions provide independent confirmation. For comparison, the bottom panel of Table 6 shows non-lattice determinations of the  $B_{(s)}$ -meson decay constants which include error estimates. These are consistent with the lattice values, but with much larger uncertainties.

The lattice-QCD results in Table 6 and Table 7 were all obtained using isospin-symmetric gauge-field configurations. The most recent calculations of  $f_{B^+}$  by the HPQCD, Fermilab/MILC, and RBC/UKQCD Collaborations [69,86,88], however, include the dominant effect of nondegenerate up- and down-quark masses by evaluating the decay constant with the valence light-quark mass fixed to the physical up-quark mass. HPQCD and RBC/UKQCD also calculate  $f_{B^0}$  by fixing the valence light-quark mass equal to the physical down-quark mass [86,88]; they find differences between the  $B^+$ - and  $B^0$ -meson decay constants of  $f_{B^0} - f_{B^+} \approx 4$  MeV and  $f_{B_s}/f_{B^+} - f_{B_s}/f_{B^0} \approx 0.025$ . Inspection of Table 6 and Table 7 shows that these differences are comparable to the error on the HPQCD 12 result for  $f_B$  [89], and to the errors on the Fermilab/MILC, HPQCD 12, and ETM results for  $f_{B_s}/f_B$  [69,89,85], none of which account for isospin breaking. Therefore, to enable comparison with experimental measurements, in this review we correct those lattice-QCD results for  $B$ -meson decay constants obtained with degenerate up and down valence quarks *a posteriori* for isospin breaking before computing our averages. For the correction factors, we use the differences obtained empirically by HPQCD in Ref. 86<sup>6</sup>

$$f_{B^+} - f_B = -1.9(5) \text{ MeV}, \quad (20)$$

$$\frac{f_{B_s}}{f_{B^+}} - \frac{f_{B_s}}{f_B} = 0.012(4), \quad (21)$$

$$f_{B^0} - f_B = 1.7(5) \text{ MeV}, \quad (22)$$

$$\frac{f_{B_s}}{f_{B^0}} - \frac{f_{B_s}}{f_B} = -0.011(4). \quad (23)$$

The isospin-breaking correction factors in Eqs. (20)–(23) are well determined because of cancellations between correlated errors in the differences.

We first average the published (2+1)-flavor lattice-QCD results for the charged and neutral  $B_{(s)}$ -meson decay constants and their ratios in Table 6 and Table 7, accounting for possibly correlated uncertainties. We treat the statistical errors as correlated between the calculations of Aoki *et al.* and RBC/UKQCD because they employ the same gauge-field

<sup>6</sup> The correlated uncertainties were provided by HPQCD via private communication.

**Table 6:** Recent theoretical determinations of  $f_{B^+}$ ,  $f_{B_s}$ , and their ratio. The upper panels show results from lattice-QCD simulations with  $(2+1+1)$  and  $(2+1)$  dynamical quark flavors, respectively. For some of the lattice-QCD results, statistical and systematic errors are quoted separately. Lattice-QCD results for  $f_B$  and  $f_{B_s}/f_B$  in the isospin-symmetric limit  $m_u = m_d$  are noted with an “\*”; they are corrected by the factors in Eq. (20) and Eq. (21), respectively, before computing the averages. Preliminary conference results noted with a “†” are not included in the averages. The bottom panel shows estimates from QCD sum rules and the light-front quark model, which are not used to obtain our preferred decay-constant values.

Reference	Method	$N_f$	$f_{B^+}$ (MeV)	$f_{B_s}$ (MeV)	$f_{B_s}/f_{B^+}$
ETM 13 [85] *,†	LQCD	2+1+1	196(9)	235(9)	1.201(25)
HPQCD 13 [86]	LQCD	2+1+1	184(4)	224(5)	1.217(8)
Average	LQCD	2+1+1	184(4)	224(5)	1.217(8)
Aoki 14 [87] *,‡	LQCD	2+1	218.8(6.5)(30.8)	263.5(4.8)(36.7)	1.193(20)(44)
RBC/UKQCD 14 [88]	LQCD	2+1	195.6(6.4)(13.3)	235.4(5.2)(11.1)	1.223(14)(70)
HPQCD 12 [89] *	LQCD	2+1	191(1)(8)	228(3)(10)	1.188(12)(13)
HPQCD 12 [89] *	LQCD	2+1	189(3)(3)*	—	—
HPQCD 11 [90]	LQCD	2+1	—	225(3)(3)	—
Fermilab/MILC 11 [69]	LQCD	2+1	196.9(5.5)(7.0)	242.0(5.1)(8.0)	1.229(13)(23)
Average	LQCD	2+1	189.9(4.2)	228.6(3.8)	1.210(15)
Our average	LQCD	Both	187.1(4.2)	227.2(3.4)	1.215(7)
Wang 15 [71] §	QCD SR		194(15)	231(16)	1.19(10)
Baker 13 [91]	QCD SR		186(14)	222(12)	1.19(4)
Lucha 13 [92]	QCD SR		192.0(14.6)	228.0(19.8)	1.184(24)
Gelhausen 13 [72]	QCD SR		207(+17/-9)	242(+17/-12)	1.17(+3/-4)
Narison 12 [73]	QCD SR		206(7)	234(5)	1.14(3)
Hwang 09 [75]	LFQM		—	270.0(42.8)¶	1.32(8)

† Lattice 2013 conference proceedings.

‡ Obtained with static  $b$  quarks (*i.e.*  $m_b \rightarrow \infty$ ).

\* Obtained by combining  $f_{B_s}$  from HPQCD 11 with  $f_{B_s}/f_B$  from this work. Approximate statistical (systematic) error obtained from quadrature sum of individual statistical (systematic) errors.

§ Obtained using  $m_b^{\text{MS}}$ ; results using  $m_b^{\text{pole}}$  are also given in the paper.

¶ Obtained by combining PDG value  $f_B = 204(31)$  MeV [76] with  $f_{B_s}/f_B$  from this work.

configurations<sup>7</sup> [87,88]. We also treat the statistical errors as correlated between the HPQCD and Fermilab/MILC calculations because they analyze an overlapping set of gauge-field configurations [69,89,90]. For  $f_{B_s}$ , we include HPQCD’s results from both 2011 [90] and 2012 [89], which were obtained using different  $b$ -quark actions, but on some of the same gauge-field configurations. HPQCD 11 and 12 also use the same determination of the absolute lattice scale, which is the second-largest

<sup>7</sup> There may be mild correlations between some sub-dominant systematic errors of Aoki *et al.* and RBC/UKQCD, who use the same determinations of the absolute lattice scale and the physical light- and strange-quark masses from Ref. 93, and who use the same power-counting estimates for the light-quark and gluon discretization errors. The effects of any correlations between these systematics, however, would be too small to impact the numerical values of the averages.

source of systematic uncertainty in both calculations. We therefore treat the statistical and scale errors as correlated between HPQCD’s  $(2+1)$ -flavor  $f_{B_s}$  results. HPQCD also presents two results for  $f_B$  in Ref. 89. The more precise value is obtained by combining the ratio  $f_{B_s}/f_B$  from this work with  $f_{B_s}$  from Ref. 90, but an associated error budget is not provided. Because this would be needed to estimate correlations between the two  $f_B$  determinations, we include only HPQCD’s more precise  $(2+1)$ -flavor result for  $f_B$  in our average. Our separate three- and four-flavor averages for the  $B^{+-}$ ,  $B^{0-}$ , and  $B_s$ -meson decay constants and ratios are listed in the lines labeled “Average” in Table 6 and Table 7, where the error on the  $(2+1)$ -flavor  $f_{B_s}$  average has been rescaled by the factor  $\sqrt{\chi^2/\text{dof}} = 1.2$  to account for the tension among results. Our  $(2+1+1)$ -flavor “averages” are identical to the “HPQCD 13” entries in Table 6

# Meson Particle Listings

 $D_s^\pm$ 

and Table 7, which are the only published four-flavor results available.

**Table 7:** Recent lattice-QCD determinations of  $f_{B^0}$  and  $f_{B_s}/f_{B^0}$ . Results obtained in the isospin-symmetric limit  $m_u = m_d$  are noted with an “\*”, while those for the  $B^+$ -meson are noted with an “ $^\S$ ”. Although the quoted results are identical to those in Table 6, they are corrected by different factors in Eq. (20)–Eq. (23) before computing the averages. Other labels and descriptions are the same as in Table 6.

Reference	Method	$N_f$	$f_{B^0}$ (MeV)	$f_{B_s}/f_{B^0}$
ETM 13 [85] $^{*,\dagger}$	LQCD	2+1+1	196(9)	1.201(25)
HPQCD 13 [86]	LQCD	2+1+1	188(4)	1.194(7)
Average	LQCD	2+1+1	188(4)	1.194(7)
Aoki 14 [87] $^{*,\ddagger}$	LQCD	2+1	218.8(6.5)(30.8)	1.193(20)(44)
RBC/UKQCD 14 [88]	LQCD	2+1	199.5(6.2)(12.6)	1.197(13)(49)
HPQCD 12 [89] *	LQCD	2+1	191(1)(8)	1.188(12)(13)
HPQCD 12 [89] *	LQCD	2+1	189(3)(3)*	–
Fermilab/MILC 11 $^\S$ [69]	LQCD	2+1	196.9(5.5)(7.0)	1.229(13)(23)
Average	LQCD	2+1	193.6(4.2)	1.187(15)
Our average	LQCD	Both	190.9(4.1)	1.192(6)

$^\dagger$  Lattice 2013 conference proceedings.

$^\ddagger$  Obtained with static  $b$  quarks (*i.e.*,  $m_b \rightarrow \infty$ ).

\* Obtained by combining  $f_{B_s}$  from HPQCD 11 with  $f_{B_s}/f_B$  from this work. Approximate statistical (systematic) error obtained from quadrature sum of individual statistical (systematic) errors.

To obtain the single-best values of the  $B_{(s)}$ -meson decay constants for phenomenology applications, we combine the available  $(2+1)$ - and  $(2+1+1)$ -flavor lattice-QCD results, which are compatible within the current level of precision. Because the four-flavor “average” is obtained from only a single result, we do not simply combine the two lines labeled “Average” in Table 6 and Table 7, which would weight the four-flavor result too heavily. Instead, we form a single average including the published  $(2+1)$ -flavor results and the  $(2+1+1)$ -flavor result from HPQCD 13. We account for the omission of charm sea-quark contributions in the three-flavor calculations by adding to the errors on the  $(2+1)$ -flavor averages in Table 6 and Table 7 our power-counting estimates of charm sea-quark errors from Sec. II.B, taking charm sea-quark error to be 100% correlated between the three-flavor results. Because the estimated charm sea-quark errors of 0.7% for decay constants and 0.2% for decay-constant ratios are much less than those on the  $(2+1)$ -flavor averages, adding them in quadrature has a tiny impact on the total uncertainties. The largest observed change is an 0.3 MeV increase on the error  $f_{B_s}$  from HPQCD 11, and most are negligible. In the combined three- and four-flavor average we also consider correlations between the results of HPQCD 12 and HPQCD 13 because, although they employ different gauge-field configurations, they both use NRQCD for the  $b$ -quark action and the bottom-light axial-vector current.<sup>8</sup> We

<sup>8</sup> HPQCD 13 uses a 1-loop radiatively improved  $b$ -quark action, whereas HPQCD 12 uses tree-level action coefficients.

take both the operator-matching and relativistic errors, which are the dominant uncertainties in the decay constants, to be correlated between the two calculations. Our final preferred theoretical values for the charged  $B^+$  and neutral  $B_{(s)}^0$ -meson decay constants and their ratio are

$$\text{Our averages : } f_{B^+} = 187.1(4.2) \text{ MeV ,}$$

$$f_{B_s} = 227.2(3.4) \text{ MeV , } \frac{f_{B_s}}{f_{B^+}} = 1.215(7) , \quad (24)$$

$$f_{B^0} = 190.9(4.1) \text{ MeV , } \frac{f_{B_s}}{f_{B^0}} = 1.192(6) . \quad (25)$$

The errors on  $f_B^+$ ,  $f_B^0$ , and  $f_{B_s}$  after combining the three- and four-flavor results are only slightly smaller than those of the separate averages due to the correlations assumed.

## V. PHENOMENOLOGICAL IMPLICATIONS

### A. $|V_{ud}|$ , $|V_{us}|$ , and status of first-row unitarity

Using the average values for  $f_{\pi^+}|V_{ud}|$ ,  $f_{K^+}|V_{us}|$ , and their ratio from Eq. (10)–Eq. (12) and for  $f_{\pi^+}$ ,  $f_{K^+}$ , and their ratio from Eq. (16), we obtain the following determinations of the CKM matrix elements  $|V_{ud}|$ ,  $|V_{us}|$ , and their ratio from leptonic decays within the standard model:

$$|V_{ud}| = 0.9764(2)(127)(10) , \quad |V_{us}| = 0.2255(3)(6)(3) ,$$

$$\frac{|V_{us}|}{|V_{ud}|} = 0.2314(2)(5)(2) , \quad (26)$$

where the errors are from the experimental branching fraction(s), the pseudoscalar decay constant(s), and radiative corrections, respectively. These results enable a precise test of the unitarity of the first row of the CKM matrix from leptonic decays alone (the contribution from  $|V_{ub}|$  is negligible). Using the values of  $|V_{ud}|$  and  $|V_{us}|$  from Eq. (26), we find

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 - 1 = 0.004(25) , \quad (27)$$

which is consistent with three-generation unitarity at the sub-percent level.

The determinations of  $|V_{ud}|$  and  $|V_{us}|$  from leptonic decays in Eq. (26) can be compared to those obtained from other processes. The result above for  $|V_{ud}|$  agrees with the determination from superallowed  $\beta$ -decay,  $|V_{ud}| = 0.97417(21)$  [94], but has an error more than fifty times larger that is primarily due to the uncertainty in the theoretical determination of  $f_{\pi^+}$ . The CKM element  $|V_{us}|$  can be determined from semileptonic  $K^+ \rightarrow \pi^0 \ell^+ \nu$  decay. Here experimental measurements provide a value for the product  $f_+^{K\pi}(0)|V_{us}|$ , where  $f_+^{K\pi}(0)$  is the form-factor at zero four-momentum transfer between the initial state kaon and the final state pion. Taking the most recent experimental determination of  $|V_{us}|f_+^{K\pi}(0) = 0.2165(4)$

Both include the same contributions to the currents at one loop, but renormalization details differ.

from Moulson [26]<sup>9</sup> and the preliminary 2015 (2+1+1)-flavor FLAG average for  $f_+(0)^{K\pi} = 0.9704(24)(22)$  [12,13]<sup>10</sup> gives  $|V_{us}| = 0.22310(74)_{\text{thy}}(41)_{\text{exp}}$  from  $K_{\ell 3}$  decay. The determinations of  $|V_{us}|$  from leptonic and semileptonic kaon decays are both quite precise (with the error from leptonic decay being about 20% smaller), but the central values differ by  $2.2\sigma$ . Finally, the combination of the ratio  $|V_{us}|/|V_{ud}|$  from leptonic decays [Eq. (26)] with  $|V_{ud}|$  from  $\beta$  decay implies an alternative determination of  $|V_{us}| = 0.2254(6)$  which agrees with the value from leptonic kaon decay, but disagrees with the  $K_{\ell 3}$ -decay result at the  $2.2\sigma$  level. Collectively, these results indicate that there is some tension between theoretical calculations and/or measurements of leptonic pion and kaon decays, semileptonic kaon decays, and superallowed  $\beta$ -decay. Although this may be due to the presence of new physics, it is also important to revisit the quoted uncertainties on both the theoretical and experimental inputs.

Finally, we combine the experimental measurements of  $f_{\pi^+}|V_{ud}|$ ,  $f_{K^+}|V_{us}|$  from leptonic pseudoscalar-meson decays in Eq. (10) and Eq. (11) with determinations of the CKM elements from other decays or unitarity to infer “experimental” values for the decay constants. Assuming that there are no significant new-physics contributions to any of the input processes, the comparison of these results with theoretical calculations of the decay constants enables a test of lattice-QCD methods. Taking  $|V_{ud}|$  from superallowed  $\beta$ -decay [100] leads to

$$f_{\pi^-}^{\text{“exp”}} = 130.50(1)(3)(13) \text{ MeV}, \quad (28)$$

where the uncertainties are from the errors on  $\Gamma$ ,  $|V_{ud}|$ , and higher-order corrections, respectively. This agrees with the theoretical value  $f_{\pi^+} = 130.2(1.7) \text{ MeV}$  in Eq. (16) obtained from an average of recent (2+1)-flavor lattice-QCD results [39,37,35]. We take the value  $|V_{us}| = 0.22534(65)$  from the most recent global unitarity-triangle fit of the UTfit Collaboration [101] because there is tension between the values of  $|V_{us}|$  obtained from leptonic and semileptonic kaon decays. This implies

$$f_{K^-}^{\text{“exp”}} = 155.72(17)(45)(16) \text{ MeV} \quad (29)$$

where the uncertainties are from the errors on  $\Gamma$ ,  $|V_{us}|$ , and higher-order corrections, respectively. This agrees with the theoretical value  $f_{K^+} = 155.6(0.4) \text{ MeV}$  in Eq. (16) obtained from an average of recent three and four-flavor lattice-QCD results [31–33,35,37,39].

<sup>9</sup> This is an update of the 2010 Flavianet review [28] that includes new measurements of the  $K_s$  lifetime [95,96],  $\text{Re}(\epsilon'/\epsilon)$  [96], and  $\mathcal{B}(K^\pm \rightarrow \pi^\pm \pi^+ \pi^-)$  [27]. The latter measurement is the primary source of the reduced error on  $\mathcal{B}(K_{\ell 3})$ , via the constraint that the sum of all branching ratios must equal unity.

<sup>10</sup> This result comes from the calculation of FNAL/MILC in Ref. 97. For comparison, the 2015 preliminary (2+1)-flavor FLAG average based on the calculations of FNAL/MILC [98] and RBC/UKQCD [99] is  $f_+(0)^{K\pi} = 0.9677(37)$ .

### B. $|V_{cd}|$ , $|V_{cs}|$ , and status of second-row unitarity

Using the average values for  $|V_{cd}|f_{D^+}$  and  $|V_{cs}|f_{D_s^+}$  from Table 2 and Table 3, and for  $f_{D^+}$  and  $f_{D_s^+}$  from Eq. (19), we obtain the following determinations of the CKM matrix elements  $|V_{cd}|$  and  $|V_{cs}|$ , and from leptonic decays within the standard model:

$$|V_{cd}| = 0.217(5)(1) \quad \text{and} \quad |V_{cs}| = 1.007(16)(5), \quad (30)$$

where the errors are from experiment and theory, respectively, and are currently limited by the measured uncertainties on the decay rates. The central value of  $|V_{cs}|$  is greater than one, but is compatible with unity within the error. The above results for  $|V_{cd}|$  and  $|V_{cs}|$  do not include higher-order electroweak and hadronic corrections to the rate, in analogy to Eq. (3). These corrections have not been computed for  $D_{(s)}^+$ -meson leptonic decays, but are estimated to be about 1–2% for charged pion and kaon decays (see Sec. II.A). Now that the uncertainties on  $|V_{cd}|$  and  $|V_{cs}|$  from leptonic decays are at this level, we hope that the needed theoretical calculations will be undertaken.

The CKM elements  $|V_{cd}|$  and  $|V_{cs}|$  can also be obtained from semileptonic  $D^+ \rightarrow \pi^0 \ell^+ \nu$  and  $D_s^+ \rightarrow K^0 \ell^+ \nu$  decays, respectively. Here experimental measurements determine the product of the form factor times the CKM element, and theory provides the value for the form factor at zero four-momentum transfer between the initial  $D_{(s)}$  meson and the final pion or kaon. We combine the latest experimental averages for  $f_+^{D\pi}(0)|V_{cd}| = 0.1425(19)$  and  $f_+^{D_s K}(0)|V_{cs}| = 0.728(5)$  from the Heavy Flavor Averaging Group (HFAG) [102] with the zero-momentum-transfer form factors  $f_+^{D\pi}(0) = 0.666(29)$  and  $f_+^{D_s K}(0) = 0.747(19)$  calculated in (2+1)-flavor lattice QCD by the HPQCD Collaboration [103,104] to obtain  $|V_{cd}| = 0.2140(97)$  and  $|V_{cs}| = 0.9746(257)$  from semileptonic  $D_{(s)}$ -meson decays. The values of  $|V_{cd}|$  from leptonic and semileptonic decays agree, while those for  $|V_{cs}|$  are compatible at the  $1.1\sigma$  level. The determinations of  $|V_{cd}|$  and  $|V_{cs}|$  from leptonic decays in Eq. (30), however, are  $2.0\times$  and  $1.6\times$  more precise than those from semileptonic decays, respectively.

The results for  $|V_{cd}|$  and  $|V_{cs}|$  from Eq. (30) enable a test of the unitarity of the second row of the CKM matrix. We obtain

$$|V_{cd}|^2 + |V_{cs}|^2 + |V_{cb}|^2 - 1 = 0.064(36), \quad (31)$$

which is in slight tension with three-generation unitarity at the  $2\sigma$  level. Because the contribution to Eq. (31) from  $|V_{cb}|$  is so small, we obtain the same result taking  $|V_{cb}|^{\text{incl.}} \times 10^3 = 42.21(78)$  from inclusive  $B \rightarrow X_c \ell \nu$  decay [105] or  $|V_{cb}|^{\text{excl.}} \times 10^3 = 39.04(75)$  from exclusive  $B \rightarrow D^* \ell \nu$  decay at zero recoil [106].

We can also combine the experimental measurements of  $f_{D^+}|V_{cd}| = 45.91(1.05) \text{ MeV}$  and  $f_{D_s^+}|V_{cs}| = 250.9(4.0) \text{ MeV}$  from leptonic pseudoscalar-meson decays from Table 2 and Table 3 with determinations of  $|V_{cd}|$  and  $|V_{cs}|$  from CKM unitarity to infer “experimental” values for the decay constants within

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the standard model. For this purpose, we obtain the values of  $|V_{cd}|$  and  $|V_{cs}|$  by relating them to other CKM elements using the Wolfenstein parameterization [107]. We take  $|V_{cd}|$  to equal the value of  $|V_{us}|$  minus the leading correction [108]:

$$|V_{cd}| = |V_{us}| \left| -1 + \frac{|V_{cb}|^2}{2}(1 - 2(\rho + i\eta)) \right| \quad (32)$$

$$= |V_{us}| \left( \left[ -1 + (1 - 2\rho) \frac{|V_{cb}|^2}{2} \right]^2 + \eta^2 |V_{cb}|^4 \right)^{1/2}. \quad (33)$$

Using  $|V_{us}| = 0.2255(3)(6)(3)$  from leptonic kaon decay, Eq. (26), inclusive  $|V_{cb}|$  as above, and  $(\rho, \eta) = (0.136(24), 0.361(14))$  from CKM unitarity [101]  $|V_{cd}| = 0.2254(7)$ . We take  $|V_{cs}| = |V_{ud}| - |V_{cb}|^2/2$  [108], using  $|V_{ud}| = 0.97417(21)$  from  $\beta$  decay [94], giving  $|V_{cs}| = 0.9733(2)$ . Given these choices, we find

$$f_{D^+}^{\text{exp}} = 203.7(4.7)(0.6) \text{ MeV} \quad \text{and} \quad f_{D_s^+}^{\text{exp}} = 257.8(4.1)(0.1) \text{ MeV}, \quad (34)$$

where the uncertainties are from the errors on  $\Gamma$  and  $|V_{us}|$  (or  $|V_{ud}|$ ), respectively. These disagree with the theoretical values  $f_{D^+} = 211.9(1.1) \text{ MeV}$  and  $f_{D_s^+} = 249.0(1.2) \text{ MeV}$  in Eq. (19) obtained from averaging recently published three and four-flavor lattice-QCD results at the  $1.7\sigma$  and  $2.0\sigma$  levels, respectively. The significances of the tensions are sensitive, however, to the choices made for  $|V_{us}|$  and  $|V_{ud}|$ . Thus resolving the inconsistencies between determinations of elements of the first row of the CKM matrix discussed previously in Sec. V.A may also reduce the mild tensions observed here.

## C. $|V_{ub}|$ and other applications

Using the average value for  $|V_{ub}|f_{B^+}$  from Table 5, and for  $f_{B^+}$  from Eq. (24), we obtain the following determination of the CKM matrix element  $|V_{ub}|$  from leptonic decays within the standard model:

$$|V_{ub}| = 4.12(37)(9) \times 10^{-3}, \quad (35)$$

where the errors are from experiment and theory, respectively. We note, however, that decays involving the third generation of quarks and leptons may be particularly sensitive to new physics associated with electroweak symmetry breaking due to their larger masses [4,6], so Eq. (35) is more likely to be influenced by new physics than the determinations of the elements of the first and second rows of the CKM matrix in the previous sections.

The CKM element  $|V_{ub}|$  can also be obtained from semileptonic  $B$ -meson decays. Over the past several years there has remained a persistent  $2\text{--}3\sigma$  tension between the determinations of  $|V_{ub}|$  from exclusive  $B \rightarrow \pi \ell \nu$  decay and from inclusive  $B \rightarrow X_u \ell \nu$  decay, where  $X_u$  denotes all hadrons which contain a constituent up quark [3,102,109–111]. The currently most precise determination of  $|V_{ub}|^{\text{excl}} = 3.72(16) \times 10^{-3}$  is obtained from a joint  $z$ -fit of the vector and scalar form

factors  $f_+^{B\pi}(q^2)$  and  $f_0^{B\pi}(q^2)$  calculated in (2+1)-flavor lattice QCD by the FNAL/MILC Collaboration [112] and experimental measurements of the differential decay rate from BaBar [113,114] and Belle [115,116]. On the other hand, the most recent PDG average of inclusive determinations obtained using the theoretical frameworks in Refs. 117–119 is  $|V_{ub}|^{\text{incl}} = 4.49(16) \left( \begin{smallmatrix} +16 \\ -18 \end{smallmatrix} \right) \times 10^{-3}$  [120]. The result for  $|V_{ub}|$  from leptonic  $B \rightarrow \tau \nu$  decay in Eq. (35) is compatible with determinations from both exclusive and inclusive semileptonic  $B$ -meson decays.

The CKM element  $|V_{ub}|$  can now also be obtained from semileptonic  $\Lambda_b$  decays. Specifically, the recent LHCb measurement of the ratio of decay rates for  $\Lambda_b \rightarrow p \ell \nu$  over  $\Lambda_b \rightarrow \Lambda_c \ell \nu$  [121], when combined with the ratio of form factors from (2+1)-flavor lattice QCD [122], enables the first determination of the ratio of CKM elements  $|V_{ub}|/|V_{cb}| = 0.083(4)(4)$  from baryonic decay. Taking  $|V_{cb}|^{\text{incl}} = 42.21(78) \times 10^{-3}$  [105] for the denominator,<sup>1</sup> we obtain  $|V_{ub}| = 3.50(17)(17)(6) \times 10^{-3}$  from exclusive  $\Lambda_b$  semileptonic decays, where the errors are from experiment, the form factors, and  $|V_{cb}|$ , respectively. The result for  $|V_{ub}|$  from leptonic  $B \rightarrow \tau \nu$  decay in Eq. (35) is  $1.4\sigma$  higher than the determination from  $b$ -baryon decays.

Given these results, the “ $V_{ub}$ ” puzzle still stands, and the determination from leptonic  $B^+$ -meson decay is not yet sufficiently precise to weigh in on the discrepancy. New and improved experimental measurements and theoretical calculations of other  $b \rightarrow u$  flavor-changing processes, however, are providing additional information and sharpening the picture of the various tensions. Further, the error on  $|V_{ub}|$  from  $B \rightarrow \tau \nu$  decay will shrink once improved rate measurements from the Belle II experiment are available.

Finally, we can combine the experimental measurement of  $|V_{ub}|f_{B^+}$  from leptonic  $B^+$ -meson decays in Table 5 with a determination of the CKM element  $|V_{ub}|$  from elsewhere to infer an “experimental” values for  $f_{B^+}$  within the standard model. This, of course, assumes that there are no significant new-physics contributions to  $B^+ \rightarrow \tau \nu$ , which may turn out not to be the case. Further, one does not know *a priori* what value to take for  $|V_{ub}|$  given the inconsistencies between the various determinations discussed above. We therefore take the PDG weighted average of the determinations from inclusive and exclusive semileptonic  $B$ -meson decays  $|V_{ub}|^{\text{excl+incl}} = 4.09(39) \times 10^{-3}$  [120], where

<sup>1</sup>This differs from the choice for  $|V_{cb}|$  made by LHCb [121], who use the determination from exclusive  $B \rightarrow D^{(*)} \ell \nu$  decays at zero recoil [123]. The Belle Experiment recently obtained a new measurement of the  $B \rightarrow D \ell \nu$  differential decay rate [124] and determination of  $|V_{cb}| = 40.83(1.13) \times 10^{-3}$ . They find that the inclusion of experimental and theoretical nonzero-recoil information increases the value for  $|V_{cb}|$  compared to when only zero-recoil information is used, and leads to agreement with the inclusive result.



the error has been rescaled by the  $\sqrt{\chi^2/\text{dof}} = 2.6$  to account for the disagreement. Using this result we obtain

$$f_{B^+}^{\text{exp}} = 188(17)(18) \text{ MeV}, \quad (36)$$

where the uncertainties are from the errors on  $\Gamma$  and  $|V_{ub}|$ , respectively. This agrees within large uncertainties with the theoretical value  $f_{B^+} = 187.1(4.2) \text{ MeV}$  in Eq. (24) obtained from an average of recent three and four-flavor lattice-QCD results [69,86,88,89].

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## References

1. K. Nakamura *et al.* (PDG), J. Phys. **G37**, 075021 (2010).
2. J. Beringer *et al.* (PDG), Phys. Rev. **D86**, 010001 (2012).
3. K.A. Olive *et al.* (PDG), Chin. Phys. C **38**, 090001 (2014).
4. W.-S. Hou, Phys. Rev. **D48**, 2342 (1993).
5. A. Akeroyd and S. Recksiegel, Phys. Lett. **B554**, 38 (2003).
6. A. Akeroyd and S. Recksiegl, J. Phys. **G29**, 2311 (2003).
7. A. Akeroyd, Prog. Theor. Phys. **111**, 295 (2004).
8. B. Dobrescu and A. Kronfeld, Phys. Rev. Lett. **100**, 241802 (2008).
9. J. Hewett, [arXiv:hep-ph/9505246](#).
10. A. Crivellin, Phys. Rev. **D81**, 031301 (2010).
11. F. Bernlochner, Z. Ligeti, and S. Turczyk, Phys. Rev. **D90**, 094003 (2014).
12. A. Jüttner, “*Light Flavour Physics*” plenary talk presented at Lattice 2015.
13. P. A. Boyle, T. Kaneko, and S. Simula (FLAG), private communication, 2015.
14. S. Aoki *et al.* (FLAG), Eur. Phys. J. **C74**, 2890 (2014).
15. W. Marciano, Phys. Rev. Lett. **93**, 231803 (2004).
16. V. Cirigliano *et al.*, Rev. Mod. Phys. **84**, 399 (2012).
17. V. Cirigliano and I. Rosell, JHEP **10**, 005 (2007).
18. A. Sirlin, Nucl. Phys. **B196**, 83 (1982).
19. T. Kinoshita, Phys. Rev. Lett. **2**, 477 (1959).
20. M. Knecht *et al.*, Eur. Phys. J. **C12**, 469 (2000).
21. B. Ananthanarayan and B. Moussallam, JHEP **06**, 047 (2004).
22. S. Descotes-Genon and B. Moussallam, Eur. Phys. J. **C42**, 403 (2005).
23. W. Marciano and A. Sirlin, Phys. Rev. Lett. **71**, 3629 (1993).
24. V. Cirigliano and H. Neufeld, Phys. Lett. **B700**, 7 (2011).
25. V. Cirigliano and I. Rosell, Phys. Rev. Lett. **99**, 231801 (2007).
26. M. Moulson, [arXiv:1411.5252](#).
27. D. Babusi *et al.* (KLOE KLOE-2), Phys. Lett. **B738**, 128 (2014).
28. M. Antonelli *et al.* (FlaviaNet Working Group on Kaon Decays), Eur. Phys. J. **C69**, 399 (2010).
29. S. Hashimoto, J. Laiho, and S.R. Sharpe, “*Lattice Quantum Chromodynamics*,” in K.A. Olive *et al.* (PDG) Chin. Phys. C **38**, 090001 (2014) and 2015 update.
30. A. Kronfeld, Ann. Rev. Nucl. and Part. Sci. **62**, 265 (2012).
31. N. Carrasco *et al.* (ETM), Phys. Rev. **D91**, 054507 (2015).
32. A. Bazavov *et al.* (Fermilab Lattice and MILC), Phys. Rev. **D90**, 074509 (2014).
33. R.J. Dowdall *et al.* (HPQCD), Phys. Rev. **D88**, 074504 (2013).
34. T. Blum *et al.* (RBC/UKQCD), [arXiv:1411.7017](#).
35. R. Arthur *et al.* (RBC/UKQCD), Phys. Rev. **D87**, 094514 (2013).
36. J. Laiho and R. Van de Water, PoS **LATTICE2011**, 293 (2011).
37. A. Bazavov *et al.* (MILC), PoS **LATTICE2010**, 074 (2010).
38. S. Durr *et al.* (BMW), Phys. Rev. **D81**, 054507 (2010).
39. E. Follana *et al.* (HPQCD, UKQCD), Phys. Rev. Lett. **100**, 062002 (2008).
40. J. Gasser and H. Leutwyler, Nucl. Phys. **B250**, 465 (1985).
41. G. de Divitiis *et al.* (RM123), JHEP **04**, 124 (2012).
42. M. Nobes, [arXiv:hep-lat/0501009](#).
43. A. Bazavov *et al.* (Fermilab Lattice and MILC), [arXiv:1602.03560](#).
44. M. Artuso *et al.* (CLEO), Phys. Rev. Lett. **95**, 251801 (2005).
45. B. Eisenstein *et al.* (CLEO), Phys. Rev. **D78**, 052003 (2008).
46. M. Ablikim *et al.* (BESIII), Phys. Rev. **D89**, 051104 (2014).
47. M. Artuso *et al.* (CLEO), Phys. Rev. Lett. **99**, 071802 (2007).
48. J. Alexander *et al.* (CLEO), Phys. Rev. **D79**, 052001 (2009).
49. A. Zupanc *et al.* (Belle), JHEP **1309**, 139 (2013).
50. P. Naik *et al.* (CLEO), Phys. Rev. **D80**, 112004 (2009).
51. K. Ecklund *et al.* (CLEO), Phys. Rev. Lett. **100**, 161801 (2008).
52. P. Oniyisi *et al.* (CLEO), Phys. Rev. **D79**, 052002 (2009).
53. P. del Amo Sanchez *et al.* (BaBar), Phys. Rev. **D82**, 091103 (2010).
54. J. Alexander *et al.* (CLEO), Phys. Rev. Lett. **100**, 161804 (2008).
55. B. Aubert *et al.* (BaBar), Phys. Rev. Lett. **98**, 141801 (2007).

## Meson Particle Listings

 $D_s^\pm$ 

56. J. Lees *et al.* (BaBar), [arXiv:1003.3063](#).
57. G. Burdman, J. Goldman, and D. Wyler, *Phys. Rev.* **D51**, 111 (1995).
58. D. Atwood, G. Eilamn, and A. Soni, *Mod. Phys. Lett.* **A11**, 1061 (1996).
59. P. Colangelo, F. De Fazio, and G. Nardulli, *Phys. Lett.* **B372**, 331 (1996).
60. A. Khodjamirian, G. Stoll, and D. Wyler, *Phys. Lett.* **B358**, 129 (1995).
61. G. Eilam, I. Halperin, and R. Mendel, *Phys. Lett.* **B361**, 137 (1995).
62. C. Geng, C. Lih and W.-M. Zhang, *Phys. Rev.* **D57**, 5697 (1998).
63. C. Geng, C. Lih and W.-M. Zhang, *Mod. Phys. Lett.* **A15**, 2087 (2000).
64. G. Korchemsky, D. Pirjol, and T.-M. Yan, *Phys. Rev.* **D61**, 114510 (2000).
65. C.-W. Hwang, *Eur. Phys. J.* **C46**, 379 (2006).
66. C.-D. Lu and G.-L. Song, *Phys. Lett.* **B562**, 75 (2003).
67. Y.-B. Yang *et al.* ( $\chi$ QCD), *Phys. Rev.* **D92**, 034517 (2015).
68. H. Na *et al.* (HPQCD), *Phys. Rev.* **D86**, 054510 (2012).
69. A. Bazavov *et al.* (Fermilab Lattice and MILC), *Phys. Rev.* **D85**, 114506 (2012).
70. C.T.H. Davies *et al.* (HPQCD), *Phys. Rev.* **D82**, 114504 (2010).
71. Z.-G. Wang, *Eur. Phys. J.* **C75**, 427 (2015).
72. P. Gelhausen *et al.*, *Phys. Rev.* **D88**, 014015 (2013).
73. S. Narison, *Phys. Lett.* **B718**, 1321 (2013).
74. W. Lucha, D. Melikhov, and S. Simula, *Phys. Lett.* **B701**, 82 (2011).
75. Chien-Wen Hwang, *Phys. Rev.* **D81**, 054022 (2010).
76. C. Amsler *et al.* (PDG), *Phys. Lett.* **B667**, 1 (2008), and 2009 update.
77. J. Laiho, E. Lunghi, and R. Van de Water, *Phys. Rev.* **D81**, 034503 (2010).
78. M. Schmelling, *Phys. Scripta* **51**, 676 (1995).
79. A. Vladikas, [arXiv:1509.01155](#).
80. K. Ikado *et al.* (Belle), *Phys. Rev. Lett.* **97**, 251802 (2006).
81. I. Adachi *et al.* (Belle), *Phys. Rev. Lett.* **110**, 131801 (2013).
82. B. Kronenbitter *et al.* (Belle), *Phys. Rev.* **D92**, 051102 (2015).
83. J. Lees *et al.* (BaBar), *Phys. Rev.* **D88**, 031102 (2013).
84. B. Aubert *et al.* (BaBar), *Phys. Rev.* **D81**, 051101 (2010).
85. N. Carrasco *et al.* (ETM), *PoS LATTICE2013*, 313 (2014).
86. R. J. Dowdall *et al.* (HPQCD), *Phys. Rev. Lett.* **110**, 222003 (2013).
87. Y. Aoki *et al.*, *Phys. Rev.* **D91**, 114505 (2015).
88. N. Christ *et al.* (RBC/UKQCD), *Phys. Rev.* **D91**, 054502 (2015).
89. H. Na *et al.* (HPQCD), *Phys. Rev.* **D86**, 034506 (2012).
90. C. McNeile *et al.* (HPQCD), *Phys. Rev.* **D85**, 031503 (2012).
91. M. Baker *et al.*, *JHEP* **07**, 032 (2014).
92. W. Lucha, D. Melikhov, and S. Simula, *Phys. Rev.* **D88**, 056011 (2013).
93. Y. Aoki *et al.* (RBC/UKQCD), *Phys. Rev.* **D83**, 074508 (2011).
94. J. Hardy and I. Towner, *Phys. Rev.* **C91**, 025501 (2015).
95. F. Ambrosino *et al.* (KLOE), *Eur. Phys. J.* **C71**, 1604 (2011).
96. E. Abouzaid *et al.* (KTeV), *Phys. Rev.* **D83**, 092001 (2011).
97. A. Bazavov *et al.* (Fermilab Lattice and MILC), *Phys. Rev. Lett.* **112**, 112001 (2014).
98. A. Bazavov *et al.* (Fermilab Lattice and MILC), *Phys. Rev.* **D87**, 073012 (2013).
99. P. Boyle *et al.* (RBC/UKQCD), *JHEP* **06**, 164 (2015).
100. I. Towner and J. Hardy, *Phys. Rev.* **C91**, 015501 (2015).
101. M. Bona *et al.* (UTfit), *JHEP* **10**, 081 (2006).
102. Y. Amhis *et al.* (HFAG), [arXiv:1412.7515](#).
103. H. Na *et al.* (HPQCD), *Phys. Rev.* **D82**, 114506 (2010).
104. H. Na *et al.* (HPQCD), *Phys. Rev.* **D84**, 114505 (2011).
105. A. Alberti *et al.*, *Phys. Rev. Lett.* **114**, 061802 (2015).
106. J. Bailey *et al.* (Fermilab Lattice, MILC), *Phys. Rev.* **D89**, 114504 (2014).
107. L. Wolfenstein, *Phys. Rev. Lett.* **51**, 1945 (1983).
108. J. Charles *et al.* (CKMfitter Group), *Eur. Phys. J.* **C41**, 1 (2005).
109. M. Antonelli *et al.*, *Phys. Rev.* **494**, 197 (2010).
110. J. Butler *et al.* (Quark Flavor Physics Working Group), [arXiv:1311.1076](#).
111. A. Bevan *et al.* (Belle, BaBar), *Eur. Phys. J.* **C74**, 3026 (2014).
112. J. Bailey *et al.* (Fermilab Lattice and MILC), *Phys. Rev.* **D92**, 014024 (2015).
113. P. del Amo Sanchez *et al.* (BaBar), *Phys. Rev.* **D83**, 032007 (2011).
114. J. Lees *et al.* (BaBar), *Phys. Rev.* **D86**, 092004 (2012).
115. H. Ha *et al.* (Belle), *Phys. Rev.* **D83**, 071101 (2011).
116. A. Sibidanov *et al.* (Belle), *Phys. Rev.* **D88**, 032005 (2013).
117. S. Bosch *et al.*, *Phys. Rev. Lett.* **93**, 221801 (2004).
118. J.R. Anderson and E. Gardi, *JHEP* **01**, 097 (2006).
119. P. Gambino *et al.*, *JHEP* **10**, 058 (2007).
120. R. Kowalewski and T. Mannel, “*Semileptonic B-meson decays and the determination of  $V_{cb}$  and  $V_{ub}$* ,” in K.A. Olive *et al.* (PDG), *Chin. Phys. C* **38**, 090001 (2014) and 2015 update.
121. R. Aaij *et al.* (LHCb), *Nat. Phys.* **11**, 743 (2015).
122. W. Detmold, C. Lehner, and S. Meinel, *Phys. Rev.* **D92**, 034503 (2015).
123. R. Kowalewski and T. Mannel, “*Semileptonic B-meson decays and the determination of  $V_{cb}$  and  $V_{ub}$* ,” in K.A. Olive *et al.* (PDG), *Chin. Phys. C* **38**, 090001 (2014).
124. R. Glattauer *et al.* (Belle), *Phys. Rev.* **D93**, 032006 (2016).

$\Gamma(e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{19}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
$<0.83 \times 10^{-4}$	90	<sup>1</sup> ZUPANC	13 BELL $e^+ e^-$ at $\Upsilon(4S), \Upsilon(5S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<2.3 \times 10^{-4}$	90	DEL-AMO-SA...10J	BABR $e^+ e^-$ , 10.58 GeV
$<1.2 \times 10^{-4}$	90	ALEXANDER 09	CLEO $e^+ e^-$ at 4170 MeV
$<1.3 \times 10^{-4}$	90	PEDLAR 07A	CLEO See ALEXANDER 09

<sup>1</sup> ZUPANC 13 also gives the limit as  $<1.0 \times 10^{-4}$  at 95% CL.

$\Gamma(\mu^+ \nu_\mu)/\Gamma_{\text{total}}$	$\Gamma_{20}/\Gamma$
See the note on "Decay Constants of Charged Pseudoscalar Mesons" above.	
VALUE (units $10^{-3}$ )	EVTS

**5.56 ± 0.25 OUR AVERAGE**

5.31 ± 0.28 ± 0.20	492 ± 26	<sup>1</sup> ZUPANC	13 BELL $e^+ e^-$ at $\Upsilon(4S), \Upsilon(5S)$
6.02 ± 0.38 ± 0.34	275 ± 17	<sup>2</sup> DEL-AMO-SA...10J	BABR $e^+ e^-$ , 10.58 GeV
5.65 ± 0.45 ± 0.17	235 ± 14	ALEXANDER 09	CLEO $e^+ e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6.44 ± 0.76 ± 0.57	169 ± 18	<sup>3</sup> WIDHALM	08 BELL See ZUPANC 13
5.94 ± 0.66 ± 0.31	88	<sup>4</sup> PEDLAR	07A CLEO See ALEXANDER 09
6.8 ± 1.1 ± 1.8	553	<sup>5</sup> HEISTER	02i ALEP Z decays

- <sup>1</sup> ZUPANC 13 uses both  $\mu^+ \nu$  and  $\tau^+ \nu$  events to get  $f_{D_s} = (255.5 \pm 4.2 \pm 5.1)$  MeV.  
<sup>2</sup> DEL-AMO-SANCHEZ 10J uses  $\mu^+ \nu_\mu$  and  $\tau^+ \nu_\tau$  events together to get  $f_{D_s} = (258.6 \pm 6.4 \pm 7.5)$  MeV.  
<sup>3</sup> WIDHALM 08 gets  $f_{D_s} = (275 \pm 16 \pm 12)$  MeV from the branching fraction.  
<sup>4</sup> PEDLAR 07A also fits  $\mu^+$  and  $\tau^+$  events together and gets an effective  $\mu^+ \nu_\mu$  branching fraction of  $(6.38 \pm 0.59 \pm 0.33) \times 10^{-3}$   
<sup>5</sup> This HEISTER 02i result is not actually an independent measurement of the absolute  $\mu^+ \nu_\mu$  branching fraction, but is in fact based on our  $\phi \pi^+$  branching fraction of 3.6 ± 0.9%, so it cannot be included in our overall fit. HEISTER 02i combines its  $D_s^+ \rightarrow \tau^+ \nu_\tau$  and  $\mu^+ \nu_\mu$  branching fractions to get  $f_{D_s} = (285 \pm 19 \pm 40)$  MeV.

$\Gamma(\mu^+ \nu_\mu)/\Gamma(\phi \pi^+)$	$\Gamma_{20}/\Gamma_{34}$
See the note on "Decay Constants of Charged Pseudoscalar Mesons" above.	
VALUE	EVTS
<hr/>	

- • • We do not use the following data for averages, fits, limits, etc. • • •
- |                       |          |                            |   |
|-----------------------|----------|----------------------------|---|
| 0.143 ± 0.018 ± 0.006 | 489 ± 55 | <sup>1</sup> AUBERT        | 07v BABR $e^+ e^- \approx \Upsilon(4S)$ |
| 0.23 ± 0.06 ± 0.04    | 18       | <sup>2</sup> ALEXANDROV 00 | BEAT $\pi^-$ nucleus, 350 GeV           |
| 0.173 ± 0.023 ± 0.035 | 182      | <sup>3</sup> CHADHA        | 98 CLE2 $e^+ e^- \approx \Upsilon(4S)$  |
| 0.245 ± 0.052 ± 0.074 | 39       | <sup>4</sup> ACOSTA        | 94 CLE2 See CHADHA 98                   |
- <sup>1</sup> AUBERT 07v gets  $f_{D_s^+} = (283 \pm 17 \pm 16)$  MeV, using  $\Gamma(D_s^+ \rightarrow \phi \pi^+)/\Gamma(\text{total}) = (4.71 \pm 0.46)\%$ .  
<sup>2</sup> ALEXANDROV 00 uses  $r_D^2/r_{D_s}^2 = 0.82 \pm 0.09$  from a lattice-gauge-theory calculation to get the relative numbers of  $D^+ \rightarrow \mu^+ \nu_\mu$  and  $D_s^+ \rightarrow \mu^+ \nu_\mu$  events. The present result leads to  $f_{D_s} = (323 \pm 44 \pm 36)$  MeV.  
<sup>3</sup> CHADHA 98 obtains  $f_{D_s} = (280 \pm 19 \pm 28 \pm 34)$  MeV from this measurement, using  $\Gamma(D_s^+ \rightarrow \phi \pi^+)/\Gamma(\text{total}) = 0.036 \pm 0.009$ .  
<sup>4</sup> ACOSTA 94 obtains  $f_{D_s} = (344 \pm 37 \pm 52 \pm 42)$  MeV from this measurement, using  $\Gamma(D_s^+ \rightarrow \phi \pi^+)/\Gamma(\text{total}) = 0.037 \pm 0.009$ .

$\Gamma(\tau^+ \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{21}/\Gamma$
See the note on "Decay Constants of Charged Pseudoscalar Mesons" above.	
VALUE (units $10^{-2}$ )	EVTS
DOCUMENT ID	TECN COMMENT

**5.55 ± 0.24 OUR AVERAGE**

5.70 ± 0.21 ± 0.31 −0.30	2.2k	<sup>1</sup> ZUPANC	13 BELL $e^+ e^-$ at $\Upsilon(4S), \Upsilon(5S)$
4.96 ± 0.37 ± 0.57	748 ± 53	<sup>2</sup> DEL-AMO-SA...10J	BABR $e^- \bar{\nu}_e \nu_\tau, \mu^- \bar{\nu}_\mu \nu_\tau$
6.42 ± 0.81 ± 0.18	126 ± 16	<sup>3</sup> ALEXANDER 09	CLEO $\tau^+ \rightarrow \pi^+ \bar{\nu}_\tau$
5.52 ± 0.57 ± 0.21	155 ± 17	<sup>3</sup> NAIK	09A CLEO $\tau^+ \rightarrow \rho^+ \bar{\nu}_\tau$
5.30 ± 0.47 ± 0.22	181 ± 16	<sup>3</sup> ONYISI	09 CLEO $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6.17 ± 0.71 ± 0.34	102	<sup>4</sup> ECKLUND	08 CLEO See ONYISI 09
8.0 ± 1.3 ± 0.4	47	<sup>4</sup> PEDLAR	07A CLEO See ALEXANDER 09
5.79 ± 0.77 ± 1.84	881	<sup>5</sup> HEISTER	02i ALEP Z decays
7.0 ± 2.1 ± 2.0	22	<sup>6</sup> ABBIENDI	01L OPAL $D_s^{*+} \rightarrow \gamma D_s^+ \text{ from } Z^0$ 's
7.4 ± 2.8 ± 2.4	16	<sup>7</sup> ACCIARRI	97F L3 $D_s^{*+} \rightarrow \gamma D_s^+ \text{ from } Z^0$ 's

- <sup>1</sup> ZUPANC 13 uses both  $\mu^+ \nu$  and  $\tau^+ \nu$  events to get  $f_{D_s} = (255.5 \pm 4.2 \pm 5.1)$  MeV.  
<sup>2</sup> DEL-AMO-SANCHEZ 10J (with a small correction; see LEES 15d) uses  $\mu^+ \nu_\mu$  and  $\tau^+ \nu_\tau$  events together to get  $f_{D_s} = (259.9 \pm 6.6 \pm 7.6)$  MeV.  
<sup>3</sup> ALEXANDER 09, NAIK 09A, and ONYISI 09 use different  $\tau$  decay modes and are independent. The three papers combined give  $f_{D_s} = (259.7 \pm 7.8 \pm 3.4)$  MeV.  
<sup>4</sup> ECKLUND 08 and PEDLAR 07A are independent: ECKLUND 08 uses  $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$  events, PEDLAR 07A uses  $\tau^+ \rightarrow \pi^+ \bar{\nu}_\tau$  events.  
<sup>5</sup> HEISTER 02i combines its  $D_s^+ \rightarrow \tau^+ \nu_\tau$  and  $\mu^+ \nu_\mu$  branching fractions to get  $f_{D_s} = (285 \pm 19 \pm 40)$  MeV.  
<sup>6</sup> This ABBIENDI 01L value gives a decay constant  $f_{D_s}$  of  $(286 \pm 44 \pm 41)$  MeV.  
<sup>7</sup> The second ACCIARRI 97F error here combines in quadrature systematic (0.016) and normalization (0.018) errors. The branching fraction gives  $f_{D_s} = (309 \pm 58 \pm 33 \pm 38)$  MeV.

$\Gamma(\tau^+ \nu_\tau)/\Gamma(\mu^+ \nu_\mu)$		$\Gamma_{21}/\Gamma_{20}$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			

- |                              |          |                      |   |
|------------------------------|----------|----------------------|---|
| 10.73 ± 0.69 ± 0.56<br>−0.53 | 2.2k/492 | <sup>1</sup> ZUPANC  | 13 BELL $e^+ e^-$ at $\Upsilon(4S), \Upsilon(5S)$ |
| 11.0 ± 1.4 ± 0.6             | 102      | <sup>2</sup> ECKLUND | 08 CLEO See ONYISI 09                             |
- <sup>1</sup> This ZUPANC 13 ratio is not independent of the separate  $\tau \nu$  and  $\mu \nu$  fractions listed above.  
<sup>2</sup> This ECKLUND 08 value also uses results from PEDLAR 07A, and it is not independent of other results in these Listings. Combined with earlier CLEO results, the decay constant  $f_{D_s}$  is  $274 \pm 10 \pm 5$  MeV.

$\Gamma(K^+ K^- e^+ \nu_e)/\Gamma(K^+ K^- \pi^+)$		$\Gamma_{22}/\Gamma_{33}$	
VALUE	DOCUMENT ID	TECN	COMMENT

- • • We do not use the following data for averages, fits, limits, etc. • • •
- |                       |                     |           |                             |
|-----------------------|---------------------|-----------|-----------------------------|
| 0.558 ± 0.007 ± 0.016 | <sup>1</sup> AUBERT | 08AN BABR | $e^+ e^-$ at $\Upsilon(4S)$ |
|-----------------------|---------------------|-----------|-----------------------------|
- <sup>1</sup> This AUBERT 08AN ratio is only for the  $K^+ K^-$  mass in the range 1.01-to-1.03 GeV in the numerator and 1.0095-to-1.0295 GeV in the denominator.

$\Gamma(\phi e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{23}/\Gamma$	
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See the end of the  $D_s^+$  Listings for measurements of  $D_s^+ \rightarrow \phi e^+ \nu_e$  form factors. Unseen decay modes of the  $\phi$  are included.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.39 ± 0.23 OUR FIT</b>	Error includes scale factor of 1.8.			
<b>2.39 ± 0.23 OUR AVERAGE</b>	Error includes scale factor of 1.8.			
2.14 ± 0.17 ± 0.08	207	HIETALA	15	Uses CLEO data
2.61 ± 0.03 ± 0.17	25k	AUBERT	08AN BABR	$e^+ e^-$ at $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.36 ± 0.23 ± 0.13	106	ECKLUND	09 CLEO	See HIETALA 15
2.29 ± 0.37 ± 0.11	45	YELTON	09 CLEO	See ECKLUND 09

$\Gamma(\phi e^+ \nu_e)/\Gamma(\phi \pi^+)$		$\Gamma_{23}/\Gamma_{34}$	
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As noted in the comment column, most of these measurements use  $\phi \mu^+ \nu_\mu$  events in addition to or instead of  $\phi e^+ \nu_e$  events.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.540 ± 0.033 ± 0.048	793	LINK	02j FOCS	Uses $\phi \mu^+ \nu_\mu$
0.54 ± 0.05 ± 0.04	367	BUTLER	94 CLE2	Uses $\phi e^+ \nu_e$ and $\phi \mu^+ \nu_\mu$
0.58 ± 0.17 ± 0.07	97	FRABETTI	93G E687	Uses $\phi \mu^+ \nu_\mu$
0.57 ± 0.15 ± 0.15	104	ALBRECHT	91 ARG	Uses $\phi e^+ \nu_e$
0.49 ± 0.10 ± 0.10 −0.14	54	ALEXANDER	90B CLEO	Uses $\phi e^+ \nu_e$ and $\phi \mu^+ \nu_\mu$

$\Gamma(\eta e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{25}/\Gamma$	
--	--	----------------------	--

Unseen decay modes of the  $\eta$  are included.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.28 ± 0.24 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>2.28 ± 0.14 ± 0.19</b>	358	HIETALA	15	Uses CLEO data
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.48 ± 0.29 ± 0.13	82	YELTON	09 CLEO	See HIETALA 15

$\Gamma(\eta e^+ \nu_e)/\Gamma(\phi e^+ \nu_e)$		$\Gamma_{25}/\Gamma_{23}$	
---	--	---------------------------	--

Unseen decay modes of the  $\eta$  and the  $\phi$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.95±0.14 OUR FIT</b>	Error includes scale factor of 1.2.			
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.24±0.12±0.15	440	<sup>1</sup> BRANDENB...	95 CLE2	See HIETALA 15
<sup>1</sup> BRANDENBURG 95 uses both $e^+$ and $\mu^+$ events and makes a phase-space adjustment to use the $\mu^+$ events as $e^+$ events.				

$\Gamma(\eta'(958) e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{26}/\Gamma$	
--	--	----------------------	--

Unseen decay modes of the  $\eta'(958)$  are included.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.68 ± 0.16 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.68 ± 0.15 ± 0.06</b>	20	HIETALA	15	Uses CLEO data
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.91 ± 0.33 ± 0.05	7.5	YELTON	09 CLEO	See HIETALA 15

$\Gamma(\eta'(958) e^+ \nu_e)/\Gamma(\phi e^+ \nu_e)$		$\Gamma_{26}/\Gamma_{23}$	
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Unseen decay modes of the resonances are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.28±0.07 OUR FIT</b>	Error includes scale factor of 1.1.			
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.43±0.11±0.07	29	<sup>1</sup> BRANDENB...	95 CLE2	See HIETALA 15
<sup>1</sup> BRANDENBURG 95 uses both $e^+$ and $\mu^+$ events and makes a phase-space adjustment to use the $\mu^+$ events as $e^+$ events.				

$[\Gamma(\eta e^+ \nu_e) + \Gamma(\eta'(958) e^+ \nu_e)]/\Gamma(\phi e^+ \nu_e)$		$\Gamma_{24}/\Gamma_{23} = (\Gamma_{25} + \Gamma_{26})/\Gamma_{23}$	
--	--	---	--

Unseen decay modes of the resonances are included.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.67 ± 0.17 ± 0.17	<sup>1</sup> BRANDENB...	95 CLE2	See HIETALA 15

<sup>1</sup> This BRANDENBURG 95 data is redundant with data in previous blocks.

# Meson Particle Listings

$D_s^\pm$

$\Gamma(\omega e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{27}/\Gamma$   
A test for  $u\bar{u}$  or  $d\bar{d}$  content in the  $D_s^+$ . Neither Cabibbo-favored nor Cabibbo-suppressed decays can contribute, and  $\omega - \phi$  mixing is an unlikely explanation for any fraction above about  $2 \times 10^{-4}$ .

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<0.20	90	MARTIN	11	CLEO $e^+ e^-$ at 4170 MeV

$\Gamma(K^0 e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{28}/\Gamma$   

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.39 \pm 0.08 \pm 0.03</math></b>	42	HIETALA	15	Uses CLEO data
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.37 \pm 0.10 \pm 0.02$	14	YELTON	09	CLEO See HIETALA 15

$\Gamma(K^*(892)^0 e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{29}/\Gamma$   
Unseen decay modes of the  $K^*(892)^0$  are included.  

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.18 \pm 0.04 \pm 0.01</math></b>	32	HIETALA	15	Uses CLEO data
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.18 \pm 0.07 \pm 0.01$	7.5	YELTON	09	CLEO See HIETALA 15

$\Gamma(f_0(980) e^+ \nu_e, f_0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{30}/\Gamma$   

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.13 \pm 0.03 \pm 0.01$	42	<sup>1</sup> HIETALA	15	Uses CLEO data
$0.20 \pm 0.03 \pm 0.01$	44	ECKLUND	09	CLEO See HIETALA 15
$0.13 \pm 0.04 \pm 0.01$	13	YELTON	09	CLEO See ECKLUND 09

<sup>1</sup>HIETALA 15 uses a tighter cut on the reconstructed  $\pi^+ \pi^-$  mass ( $\pm 60$  MeV around the  $r^0$ ) than ECKLUND 09. It finds that applying the same tight cut to both analyses gives consistent results.

## Hadronic modes with a $K\bar{K}$ pair

$\Gamma(K^+ K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{31}/\Gamma$   

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.50 \pm 0.05</math> OUR FIT</b>			
<b><math>1.52 \pm 0.05 \pm 0.03</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.49 \pm 0.07 \pm 0.05$	<sup>1</sup> ALEXANDER	08	CLEO See ONYISI 13

<sup>1</sup>ALEXANDER 08 uses single- and double-tagged events in an overall fit.

$\Gamma(K^+ \bar{K}^0)/\Gamma_{\text{total}}$   $\Gamma_{32}/\Gamma$   

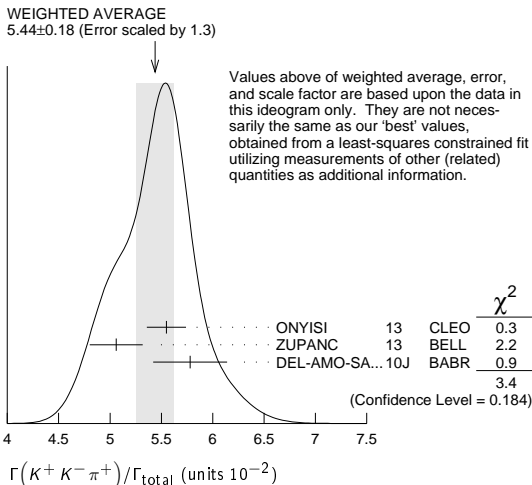
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.95 \pm 0.11 \pm 0.09</math></b>	2.0k	<sup>1</sup> ZUPANC	13	BELL $e^+ e^-$ at $\Upsilon(4S), \Upsilon(5S)$

<sup>1</sup>ZUPANC 13 finds the  $\bar{K}^0$  from its missing-mass squared, not from  $K_S^0 \rightarrow \pi^+ \pi^-$ .  
The DCS ( $D_s^+ \rightarrow K^+ K^0$ ) contribution to this fraction is estimated to be an order of magnitude below the statistical uncertainty.

$\Gamma(K^+ K^- \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{33}/\Gamma$   

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.45 \pm 0.17</math> OUR FIT</b>				Error includes scale factor of 1.2.
<b><math>5.44 \pm 0.18</math> OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
$5.55 \pm 0.14 \pm 0.13$		ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV
$5.06 \pm 0.15 \pm 0.21$	4.1k	ZUPANC	13	BELL $e^+ e^-$ at $\Upsilon(4S), \Upsilon(5S)$
$5.78 \pm 0.20 \pm 0.30$		DEL-AMO-SA...10J	BABR	$e^+ e^-$ , 10.58 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$5.50 \pm 0.23 \pm 0.16$		<sup>1</sup> ALEXANDER	08	CLEO See ONYISI 13

<sup>1</sup>ALEXANDER 08 uses single- and double-tagged events in an overall fit.



$\Gamma(\phi \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{34}/\Gamma$   
The results here are model-independent. For earlier, model-dependent results, see our PDG 06 edition. We decouple the  $D_s^+ \rightarrow \phi \pi^+$  branching fraction obtained from mass projections (and used to get some of the other branching fractions) from the  $D_s^+ \rightarrow \phi \pi^+, \phi \rightarrow K^+ K^-$  branching fraction obtained from the Dalitz-plot analysis of  $D_s^+ \rightarrow K^+ K^- \pi^+$ . That is, the ratio of these two branching fractions is not exactly the  $\phi \rightarrow K^+ K^-$  branching fraction 0.491.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.5 \pm 0.4</math> OUR AVERAGE</b>				
$4.62 \pm 0.36 \pm 0.51$		<sup>1</sup> AUBERT	06N BABR	$e^+ e^-$ at $\Upsilon(4S)$
$4.81 \pm 0.52 \pm 0.38$	212 ± 19	<sup>2</sup> AUBERT	05v BABR	$e^+ e^- \approx \Upsilon(4S)$
$3.59 \pm 0.77 \pm 0.48$		<sup>3</sup> ARTUSO	96 CLE2	$e^+ e^-$ at $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.9 \pm 5.1 \pm 1.8$		<sup>4</sup> BAI	95c BES	$e^+ e^-$ 4.03 GeV

<sup>1</sup>This AUBERT 06N measurement uses  $\bar{B}^0 \rightarrow D_s^{(*)-} D^{(*)+}$  and  $B^- \rightarrow D_s^{(*)-} D^{(*)0}$  decays, including some from other papers. However, the result is independent of AUBERT 05v.  
<sup>2</sup>AUBERT 05v uses the ratio of  $B^0 \rightarrow D^{*-} D_s^{*+}$  events seen in two different ways, in both of which the  $D^{*-} \rightarrow \bar{D}^0 \pi^-$  decay is fully reconstructed: (1) The  $D_s^{*+} \rightarrow D_s^+ \gamma$ ,  $D_s^+ \rightarrow \phi \pi^+$  decay is fully reconstructed. (2) The number of events in the  $D_s^+$  peak in the missing mass spectrum against the  $D^{*-} \gamma$  is measured.  
<sup>3</sup>ARTUSO 96 uses partially reconstructed  $\bar{B}^0 \rightarrow D^{*+} D_s^{*-}$  decays to get a model-independent value for  $\Gamma(D_s^- \rightarrow \phi \pi^-)/\Gamma(D^0 \rightarrow K^- \pi^+)$  of  $0.92 \pm 0.20 \pm 0.11$ .  
<sup>4</sup>BAI 95c uses  $e^+ e^- \rightarrow D_s^+ D_s^-$  events in which one or both of the  $D_s^\pm$  are observed to obtain the first model-independent measurement of the  $D_s^+ \rightarrow \phi \pi^+$  branching fraction, without assumptions about  $\sigma(D_s^\pm)$ . However, with only two “doubly-tagged” events, the statistical error is very large.

$\Gamma(\phi \pi^+, \phi \rightarrow K^+ K^-)/\Gamma(K^+ K^- \pi^+)$   $\Gamma_{35}/\Gamma_{33}$   
This is the “fit fraction” from the Dalitz-plot analysis. We decouple the  $D^+ \rightarrow \phi \pi^+$  branching fraction obtained from mass projections (and used to get some of the other branching fractions) from the  $D_s^+ \rightarrow \phi \pi^+, \phi \rightarrow K^+ K^-$  branching fraction obtained from the Dalitz-plot analysis of  $D_s^+ \rightarrow K^+ K^- \pi^+$ . That is, the ratio of these two branching fractions is not exactly the  $\phi \rightarrow K^+ K^-$  branching fraction 0.491.

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>41.6 \pm 0.8</math> OUR AVERAGE</b>			
$41.4 \pm 0.8 \pm 0.5$	DEL-AMO-SA...11G	BABR	Dalitz fit, 96k ± 369 evts
$42.2 \pm 1.6 \pm 0.3$	MITCHELL	09A CLEO	Dalitz fit, 12k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$39.6 \pm 3.3 \pm 4.7$	FRABETTI	95B E687	Dalitz fit, 701 evts

$\Gamma(K^+ \bar{K}^*(892)^0, \bar{K}^{*0} \rightarrow K^- \pi^+)/\Gamma(K^+ K^- \pi^+)$   $\Gamma_{36}/\Gamma_{33}$   
This is the “fit fraction” from the Dalitz-plot analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>47.8 \pm 0.6</math> OUR AVERAGE</b>			
$47.9 \pm 0.5 \pm 0.5$	DEL-AMO-SA...11G	BABR	Dalitz fit, 96k ± 369 evts
$47.4 \pm 1.5 \pm 0.4$	MITCHELL	09A CLEO	Dalitz fit, 12k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$47.8 \pm 4.6 \pm 4.0$	FRABETTI	95B E687	Dalitz fit, 701 evts

$\Gamma(f_0(980) \pi^+, f_0 \rightarrow K^+ K^-)/\Gamma(K^+ K^- \pi^+)$   $\Gamma_{37}/\Gamma_{33}$   
This is the “fit fraction” from the Dalitz-plot analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>21 \pm 6</math> OUR AVERAGE</b>			Error includes scale factor of 3.5.
$16.4 \pm 0.7 \pm 2.0$	DEL-AMO-SA...11G	BABR	Dalitz fit, 96k ± 369 evts
$28.2 \pm 1.9 \pm 1.8$	MITCHELL	09A CLEO	Dalitz fit, 12k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$11.0 \pm 3.5 \pm 2.6$	FRABETTI	95B E687	Dalitz fit, 701 evts

$\Gamma(f_0(1370) \pi^+, f_0 \rightarrow K^+ K^-)/\Gamma(K^+ K^- \pi^+)$   $\Gamma_{38}/\Gamma_{33}$   
This is the “fit fraction” from the Dalitz-plot analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>1.3 \pm 0.8</math> OUR AVERAGE</b>			Error includes scale factor of 3.9.
$1.1 \pm 0.1 \pm 0.2$	DEL-AMO-SA...11G	BABR	Dalitz fit, 96k ± 369 evts
$4.3 \pm 0.6 \pm 0.5$	MITCHELL	09A CLEO	Dalitz fit, 12k evts

$\Gamma(f_0(1710) \pi^+, f_0 \rightarrow K^+ K^-)/\Gamma(K^+ K^- \pi^+)$   $\Gamma_{39}/\Gamma_{33}$   
This is the “fit fraction” from the Dalitz-plot analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>1.2 \pm 0.5</math> OUR AVERAGE</b>			Error includes scale factor of 3.8.
$1.1 \pm 0.1 \pm 0.1$	DEL-AMO-SA...11G	BABR	Dalitz fit, 96k ± 369 evts
$3.4 \pm 0.5 \pm 0.3$	MITCHELL	09A CLEO	Dalitz fit, 12k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.4 \pm 2.3 \pm 3.5$	FRABETTI	95B E687	Dalitz fit, 701 evts

$\Gamma(K^+ \bar{K}_S^0(1430)^0, \bar{K}_S^0 \rightarrow K^- \pi^+)/\Gamma(K^+ K^- \pi^+)$   $\Gamma_{40}/\Gamma_{33}$   
This is the “fit fraction” from the Dalitz-plot analysis.  

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>3.4 \pm 0.7</math> OUR AVERAGE</b>			Error includes scale factor of 1.2.
$2.4 \pm 0.3 \pm 1.0$	DEL-AMO-SA...11G	BABR	Dalitz fit, 96k ± 369 evts
$3.9 \pm 0.5 \pm 0.5$	MITCHELL	09A CLEO	Dalitz fit, 12k evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$9.3 \pm 3.2 \pm 3.2$	FRABETTI	95B E687	Dalitz fit, 701 evts

$\Gamma(K^+ K_S^0 \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{41}/\Gamma$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.52 \pm 0.09 \pm 0.20</math></b>	ONYISI	13	CLEO	$e^+ e^-$ at 4.17 GeV

$\Gamma(2K_S^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{42}/\Gamma$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.77 \pm 0.05 \pm 0.03</math></b>	ONYISI	13	CLEO	$e^+ e^-$ at 4.17 GeV

$\Gamma(K^*(892)+\bar{K}^0)/\Gamma(\phi\pi^+)$	$\Gamma_{44}/\Gamma_{34}$			
Unseen decay modes of the resonances are included.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.20\pm0.21\pm0.13</math></b>	CHEN	89	CLEO	$e^+e^-$ 10 GeV

$\Gamma(K^+ K^- \pi^+ \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{45}/\Gamma$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>6.3 \pm 0.6</math> OUR FIT</b>				
<b><math>6.37 \pm 0.21 \pm 0.56</math></b>	ONYISI	13	CLEO	$e^+ e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$5.65 \pm 0.29 \pm 0.40$	<sup>1</sup> ALEXANDER	08	CLEO	See ONYISI 13
<sup>1</sup> ALEXANDER 08 uses single- and double-tagged events in an overall fit.				

$\Gamma(\phi \rho^+)/\Gamma(\phi \pi^+)$	$\Gamma_{46}/\Gamma_{34}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.86 \pm 0.26 \pm 0.29</math></b>	253	AVERY	92	CLE2 $e^+ e^- \simeq 10.5$ GeV

$\Gamma(K_S^0 K^- 2\pi^+)/\Gamma_{\text{total}}$	$\Gamma_{47}/\Gamma$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.67 \pm 0.10</math> OUR FIT</b>				
<b><math>1.69 \pm 0.07 \pm 0.08</math></b>	ONYISI	13	CLEO	$e^+ e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.64 \pm 0.10 \pm 0.07$	<sup>1</sup> ALEXANDER	08	CLEO	See ONYISI 13
<sup>1</sup> ALEXANDER 08 uses single- and double-tagged events in an overall fit.				

$\Gamma(K^*(892)+\bar{K}^*(892)^0)/\Gamma(\phi\pi^+)$				$\Gamma_{48}/\Gamma_{34}$
Unseen decay modes of the resonances are included.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.6\pm0.4\pm0.4</math></b>	ALBRECHT	92B	ARG	$e^+e^-\simeq 10.4$ GeV

$\Gamma(K^+ K_S^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{49}/\Gamma$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.03 \pm 0.06 \pm 0.08</math></b>	ONYISI	13	CLEO	$e^+ e^-$ at 4.17 GeV

$\Gamma(K^+ K_S^0 \pi^+ \pi^-)/\Gamma(K_S^0 K^- 2\pi^+)$	$\Gamma_{49}/\Gamma_{47}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.586 \pm 0.052 \pm 0.043</math></b>	476	LINK	01c	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV

$\Gamma(K^+ K^- 2\pi^+ \pi^-)/\Gamma(K^+ K^- \pi^+)$	$\Gamma_{50}/\Gamma_{33}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.160 \pm 0.027</math> OUR AVERAGE</b>				
$0.150 \pm 0.019 \pm 0.025$	240	LINK	03d	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV
$0.188 \pm 0.036 \pm 0.040$	75	FRABETTI	97c	E687 $\gamma Be, \bar{E}_\gamma \approx 200$ GeV

$\Gamma(\phi 2\pi^+ \pi^-)/\Gamma(\phi \pi^+)$	$\Gamma_{51}/\Gamma_{34}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.269 \pm 0.027</math> OUR AVERAGE</b>				
$0.249 \pm 0.024 \pm 0.021$	136	LINK	03d	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV
$0.28 \pm 0.06 \pm 0.01$	40	FRABETTI	97c	E687 $\gamma Be, \bar{E}_\gamma \approx 200$ GeV
$0.58 \pm 0.21 \pm 0.10$	21	FRABETTI	92	E687 $\gamma Be$
$0.42 \pm 0.13 \pm 0.07$	19	ANJOS	88	E691 Photoproduction
$1.11 \pm 0.37 \pm 0.28$	62	ALBRECHT	85d	ARG $e^+ e^- 10$ GeV

$\Gamma(K^+ K^- \rho^0 \pi^+ \text{non-}\phi)/\Gamma(K^+ K^- 2\pi^+ \pi^-)$	$\Gamma_{52}/\Gamma_{50}$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.03</math></b>	90	LINK	03d	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV

$\Gamma(\phi \rho^+ \pi^+, \phi \rightarrow K^+ K^-)/\Gamma(K^+ K^- 2\pi^+ \pi^-)$	$\Gamma_{53}/\Gamma_{50}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.75 \pm 0.06 \pm 0.04</math></b>	LINK	03d	FOCS	$\gamma A, \bar{E}_\gamma \approx 180$ GeV

$\Gamma(\phi a_1(1260)^+, \phi \rightarrow K^+ K^-, a_1^+ \rightarrow \rho^0 \pi^+)/\Gamma(K^+ K^- \pi^+)$	$\Gamma_{54}/\Gamma_{33}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.137 \pm 0.019 \pm 0.011</math></b>	LINK	03d	FOCS	$\gamma A, \bar{E}_\gamma \approx 180$ GeV

$\Gamma(K^+ K^- 2\pi^+ \pi^- \text{nonresonant})/\Gamma(K^+ K^- 2\pi^+ \pi^-)$	$\Gamma_{55}/\Gamma_{50}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.10 \pm 0.06 \pm 0.05</math></b>	LINK	03d	FOCS	$\gamma A, \bar{E}_\gamma \approx 180$ GeV

$\Gamma(2K_S^0 2\pi^+ \pi^-)/\Gamma(K_S^0 K^- 2\pi^+)$	$\Gamma_{56}/\Gamma_{47}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.051 \pm 0.015 \pm 0.015</math></b>	37 $\pm$ 10	LINK	04d	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV

$\Gamma(\pi^+ \pi^0)/\Gamma(K^+ K_S^0)$	$\Gamma_{57}/\Gamma_{31}$			
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.3</math></b>	90	MENDEZ	10	CLEO $e^+ e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4.1$	90	ADAMS	07A	CLEO See MENDEZ 10

$\Gamma(2\pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{58}/\Gamma$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.09 \pm 0.05</math> OUR FIT</b>	Error includes scale factor of 1.1.			
<b><math>1.11 \pm 0.04 \pm 0.04</math></b>	ONYISI	13	CLEO	$e^+ e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.11 \pm 0.07 \pm 0.04$	<sup>1</sup> ALEXANDER	08	CLEO	See ONYISI 13
<sup>1</sup> ALEXANDER 08 uses single- and double-tagged events in an overall fit.				

$\Gamma(2\pi^+ \pi^-)/\Gamma(K^+ K^- \pi^+)$	$\Gamma_{58}/\Gamma_{33}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.201 \pm 0.007</math> OUR FIT</b>				
<b><math>0.199 \pm 0.004 \pm 0.009</math></b>	$\approx 10.5k$	AUBERT	09o	BABR $e^+ e^- \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.265 \pm 0.041 \pm 0.031$	98	FRABETTI	97d	E687 $\gamma Be \approx 200$ GeV

$\Gamma(\rho^0 \pi^+)/\Gamma(2\pi^+ \pi^-)$	$\Gamma_{59}/\Gamma_{58}$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.018 \pm 0.005 \pm 0.010</math></b>		AUBERT	09o	BABR Dalitz fit, $\approx 10.5k$ evts
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen		LINK	04	FOCS Dalitz fit, 1475 $\pm$ 50 evts
$0.058 \pm 0.023 \pm 0.037$		AITALA	01A	E791 Dalitz fit, 848 evts
$<0.073$	90	FRABETTI	97d	E687 $\gamma Be \approx 200$ GeV

$\Gamma(\pi^+(\pi^+\pi^-)_{S\text{-wave}})/\Gamma(2\pi^+\pi^-)$	$\Gamma_{60}/\Gamma_{58}$			
<p>This is the “fit fraction” from the Dalitz-plot analysis. See also KLEMPOT 08, which uses 568 <math>D_s^+ \rightarrow 3\pi</math> decays (over 280 background events) from FNAL E791 to study various parametrizations of the decay amplitudes. The emphasis there is more on S-wave <math>\pi\pi</math> decay products — 20 different solutions are given — than on <math>D_s^+</math> fit fractions.</p>				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.833 \pm 0.020</math> OUR AVERAGE</b>				
$0.830 \pm 0.009 \pm 0.019$	<sup>1</sup> AUBERT	09o	BABR	Dalitz fit, $\approx 10.5k$ evts
$0.8704 \pm 0.0560 \pm 0.0438$	<sup>2</sup> LINK	04	FOCS	Dalitz fit, 1475 $\pm$ 50 evts

<sup>1</sup> AUBERT 09o gives the amplitude and phase of the  $\pi^+ \pi^-$  S-wave in 29  $\pi^+ \pi^-$  invariant-mass bins.

<sup>2</sup> LINK 04 borrows a K-matrix parametrization from ANISOVICH 03 of the full  $\pi\pi$  S-wave isoscalar scattering amplitude to describe the  $\pi^+ \pi^-$  S-wave component of the  $\pi^+ \pi^+ \pi^-$  state. The fit fraction given above is a sum over five  $f_0$  mesons, the  $f_0(980)$ ,  $f_0(1300)$ ,  $f_0(1200\text{--}1600)$ ,  $f_0(1500)$ , and  $f_0(1750)$ . See LINK 04 for details and discussion.

LINK 04 borrows a K-matrix parametrization from ANISOVICH 03 of the full  $\pi^+\pi^-\pi^0$  S-wave isoscalar scattering amplitude to describe the  $\pi^+\pi^-\pi^0$  S-wave component of the  $\pi^+\pi^+\pi^-\pi^0$  state. The fit fraction given above is a sum over five  $f_0$  mesons, the  $f_0(980)$ ,  $f_0(1300)$ ,  $f_0(1200-1600)$ ,  $f_0(1500)$ , and  $f_0(1750)$ . See LINK 04 for details and discussion.

$\Gamma(f_0(980)\pi^+, f_0 \rightarrow \pi^+\pi^-)/\Gamma(2\pi^+\pi^-)$				$\Gamma_{61}/\Gamma_{58}$
This is the “fit fraction” from the Dalitz-plot analysis. See above for the full $\pi^+(\pi^+\pi^-)_{S\text{-wave}}$ fit fraction.				
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.565 \pm 0.043 \pm 0.047$	AITALA	01A	E791	Dalitz fit, 848 evts
$1.074 \pm 0.140 \pm 0.043$	FRABETTI	97D	E687	$\gamma$ Be $\approx$ 200 GeV

$\Gamma(f_0(1370)\pi^+, f_0 \rightarrow \pi^+\pi^-)/\Gamma(2\pi^+\pi^-)$	$\Gamma_{62}/\Gamma_{58}$			
This is the “fit fraction” from the Dalitz-plot analysis. See above for the full $\pi^+(\pi^+\pi^-)_{S-wave}$ fit fraction.				
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.274 \pm 0.114 \pm 0.019$	<sup>1</sup> FRABETTI	97d	E687	$\gamma Be \approx 200$ GeV

<sup>1</sup> FRABETTI 97d calls this mode  $S(1475) \pi^+$ , but finds the mass and width of this  $S(1475)$  to be in excellent agreement with those of the  $f_0(1500)$ .

$\Gamma(f_0(1500)\pi^+, f_0 \rightarrow \pi^+\pi^-)/\Gamma(2\pi^+\pi^-)$	$\Gamma_{63}/\Gamma_{58}$			
This is the “fit fraction” from the Dalitz-plot analysis. See above for the full $\pi^+(\pi^+\pi^-\pi^-)_{S-wave}$ fit fraction.				
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.274 \pm 0.114 \pm 0.019$	<sup>1</sup> FRABETTI	97d	E687	$\gamma Be \approx 200$ GeV
<sup>1</sup> FRABETTI 97d calls this mode $S(1475)\pi^+$ , but finds the mass and width of this $S(1475)$ to be in excellent agreement with those of the $f_0(1500)$ .				

# Meson Particle Listings

$D_s^\pm$

## $\Gamma(\rho(1450)^0\pi^+,\rho^0\rightarrow\pi^+\pi^-)/\Gamma(2\pi^+\pi^-)$ $\Gamma_{65}/\Gamma_{58}$

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.027 ±0.018 OUR AVERAGE</b>			
0.023 ±0.008 ±0.017	AUBERT	09o	BABR Dalitz fit, ≈ 10.5k evts
0.0656±0.0343±0.0440	LINK	04	FOCS Dalitz fit, 1475 ± 50 evts
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.044 ±0.021 ±0.002	AITALA	01A	E791 Dalitz fit, 848 evts

$\Gamma(\pi^+2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{66}/\Gamma$
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>0.65±0.13±0.03</b>	72 ± 16
NAIK	09A CLEO
e <sup>+</sup> e <sup>-</sup> at 4170 MeV	

$\Gamma(2\pi^+\pi^-\pi^0)/\Gamma(\phi\pi^+)$	$\Gamma_{67}/\Gamma_{34}$
VALUE	CL%
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<3.3	90
ANJOS	89E E691
Photoproduction	

$\Gamma(\eta\pi^+)/\Gamma_{\text{total}}$	$\Gamma_{68}/\Gamma$
Unseen decay modes of the $\eta$ are included.	
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>1.70±0.09 OUR FIT</b>	Error includes scale factor of 1.1.
<b>1.71±0.08 OUR AVERAGE</b>	
1.67±0.08±0.06	ONYISI 13 CLEO
1.82±0.14±0.07	0.8k ZUPANC 13 BELL
• • • We do not use the following data for averages, fits, limits, etc. • • •	
1.58±0.11±0.18	<sup>1</sup> ALEXANDER 08 CLEO
See ONYISI 13	
<sup>1</sup> ALEXANDER 08 uses single- and double-tagged events in an overall fit.	

$\Gamma(\eta\pi^+)/\Gamma(K^+K_S^0)$	$\Gamma_{68}/\Gamma_{31}$
Unseen decay modes of the $\eta$ are included.	
VALUE	EVTs
<b>1.13 ±0.07 OUR FIT</b>	Error includes scale factor of 1.1.
• • • We do not use the following data for averages, fits, limits, etc. • • •	
1.236±0.043±0.063	2587 ± 89
MENDEZ	10 CLEO
See ONYISI 13	

$\Gamma(\eta\pi^+)/\Gamma(\phi\pi^+)$	$\Gamma_{68}/\Gamma_{34}$
Unseen decay modes of the resonances are included.	
VALUE	EVTs
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.48±0.03±0.04	920
JESSOP	98 CLE2
0.54±0.09±0.06	165
ALEXANDER	92 CLE2
See JESSOP 98	

$\Gamma(\omega\pi^+)/\Gamma_{\text{total}}$	$\Gamma_{69}/\Gamma$
Unseen decay modes of the $\omega$ are included.	
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>0.24±0.06 OUR FIT</b>	
<b>0.21±0.09±0.01</b>	6 ± 2.4
GE	09A CLEO
e <sup>+</sup> e <sup>-</sup> at 4170 MeV	

$\Gamma(\omega\pi^+)/\Gamma(\eta\pi^+)$	$\Gamma_{69}/\Gamma_{68}$
Unseen decay modes of the resonances are included.	
VALUE	DOCUMENT ID
<b>0.14±0.04 OUR FIT</b>	
<b>0.16±0.04±0.03</b>	BALEST 97 CLE2
e <sup>+</sup> e <sup>-</sup> ≈ $\Upsilon(4S)$	

$\Gamma(3\pi^+2\pi^-)/\Gamma(K^+K^-\pi^+)$	$\Gamma_{70}/\Gamma_{33}$
VALUE	EVTs
<b>0.146±0.014 OUR AVERAGE</b>	
0.145±0.011±0.010	671
LINK	03D FOCS
0.158±0.042±0.031	37
FRABETTI	97C E687
$\gamma$ A, $\overline{E}_\gamma$ ≈ 180 GeV	
$\gamma$ Be, $\overline{E}_\gamma$ ≈ 200 GeV	

$\Gamma(\eta\rho^+)/\Gamma_{\text{total}}$	$\Gamma_{72}/\Gamma$
Unseen decay modes of the $\eta$ are included.	
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>8.9±0.6±0.5</b>	328 ± 22
NAIK	09A CLEO
$\eta \rightarrow 2\gamma$	

$\Gamma(\eta\rho^+)/\Gamma(\phi\pi^+)$	$\Gamma_{72}/\Gamma_{34}$
Unseen decay modes of the resonances are included.	
VALUE	EVTs
• • • We do not use the following data for averages, fits, limits, etc. • • •	
2.98±0.20±0.39	447
JESSOP	98 CLE2
2.86±0.38 ±0.36	217
AVERY	92 CLE2
See JESSOP 98	

$\Gamma(\eta\pi^+\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{73}/\Gamma$
VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID
<b>9.2±0.4±1.1</b>	ONYISI 13 CLEO
e <sup>+</sup> e <sup>-</sup> at 4.17 GeV	

$\Gamma(\omega\pi^+\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{74}/\Gamma$
Unseen decay modes of the $\omega$ are included.	
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>2.78±0.65 ±0.25</b>	34 ± 7.9
GE	09A CLEO
e <sup>+</sup> e <sup>-</sup> at 4170 MeV	

$\Gamma(3\pi^+2\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{75}/\Gamma$
VALUE	DOCUMENT ID
<b>0.049 +0.033 -0.030</b>	BARLAG 92c ACCM
$\pi^-$ 230 GeV	

$\Gamma(\omega 2\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{76}/\Gamma$
Unseen decay modes of the $\omega$ are included.	
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>1.58±0.45 ±0.09</b>	29 ± 8.2
GE	09A CLEO
e <sup>+</sup> e <sup>-</sup> at 4170 MeV	

$\Gamma(\eta'(958)\pi^+)/\Gamma_{\text{total}}$	$\Gamma_{77}/\Gamma$
Unseen decay modes of the $\eta'(958)$ are included.	
VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID
<b>3.94±0.15 ±0.20</b>	ONYISI 13 CLEO
• • • We do not use the following data for averages, fits, limits, etc. • • •	
3.77±0.25 ±0.30	<sup>1</sup> ALEXANDER 08 CLEO
See ONYISI 13	
<sup>1</sup> ALEXANDER 08 uses single- and double-tagged events in an overall fit.	

$\Gamma(\eta'(958)\pi^+)/\Gamma(K^+K_S^0)$	$\Gamma_{77}/\Gamma_{31}$
Unseen decay modes of the $\eta'(958)$ are included.	
VALUE	EVTs
• • • We do not use the following data for averages, fits, limits, etc. • • •	
2.654±0.088±0.139	1436 ± 47
MENDEZ	10 CLEO
See ONYISI 13	

$\Gamma(\eta'(958)\pi^+)/\Gamma(\phi\pi^+)$	$\Gamma_{77}/\Gamma_{34}$
Unseen decay modes of the resonances are included.	
VALUE	EVTs
• • • We do not use the following data for averages, fits, limits, etc. • • •	
1.03±0.06 ±0.07	537
JESSOP	98 CLE2
1.20±0.15 ±0.11	281
ALEXANDER	92 CLE2
2.5 ±1.0 +1.5 -0.4	22
ALVAREZ	91 NA14
Photoproduction	
2.5 ±0.5 ±0.3	215
ALBRECHT	90D ARG
e <sup>+</sup> e <sup>-</sup> ≈ 10.4 GeV	

$\Gamma(\omega\eta\pi^+)/\Gamma_{\text{total}}$	$\Gamma_{79}/\Gamma$
Unseen decay modes of the $\omega$ and $\eta$ are included.	
VALUE	CL%
<b>&lt;2.13 × 10<sup>-2</sup></b>	90
GE	09A CLEO
e <sup>+</sup> e <sup>-</sup> at 4170 MeV	

$\Gamma(\eta'(958)\rho^+)/\Gamma_{\text{total}}$	$\Gamma_{80}/\Gamma$
VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID
<b>5.8±1.4±0.4</b>	ABLIKIM 15z BES3
482 pb <sup>-1</sup> , 4009 MeV	

$\Gamma(\eta'(958)\rho^+)/\Gamma(\phi\pi^+)$	$\Gamma_{80}/\Gamma_{34}$
Unseen decay modes of the resonances are included.	
VALUE	EVTs
<b>2.78±0.28 ±0.30</b>	137
JESSOP	98 CLE2
• • • We do not use the following data for averages, fits, limits, etc. • • •	
3.44±0.62 +0.44 -0.46	68
AVERY	92 CLE2
See JESSOP 98	

$\Gamma(\eta'(958)\pi^+\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{81}/\Gamma$
VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID
<b>5.6±0.5±0.6</b>	ONYISI 13 CLEO
e <sup>+</sup> e <sup>-</sup> at 4.17 GeV	

$\Gamma(\eta'(958)\pi^+\pi^0\text{nonresonant})/\Gamma_{\text{total}}$	$\Gamma_{82}/\Gamma$
VALUE	CL%
<b>&lt;5.1 × 10<sup>-2</sup></b>	90
ABLIKIM	15z BES3
482 pb <sup>-1</sup> , 4009 MeV	

### Modes with one or three K's

$\Gamma(K^+\pi^0)/\Gamma(K^+K_S^0)$	$\Gamma_{83}/\Gamma_{31}$
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>4.2±1.4±0.2</b>	202 ± 70
MENDEZ	10 CLEO
• • • We do not use the following data for averages, fits, limits, etc. • • •	
5.5±1.3±0.7	141 ± 34
ADAMS	07A CLEO
See MENDEZ 10	

$\Gamma(K_S^0\pi^+)/\Gamma(K^+K_S^0)$	$\Gamma_{84}/\Gamma_{31}$
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>8.12±0.28 OUR AVERAGE</b>	
8.5 ±0.7 ±0.2	393 ± 33
MENDEZ	10 CLEO
8.03±0.24±0.19	17.6k±481
WON	09 BELL
10.4 ±2.4 ±1.4	113 ± 26
LINK	08 FOCS
• • • We do not use the following data for averages, fits, limits, etc. • • •	
8.2 ±0.9 ±0.2	206 ± 22
ADAMS	07A CLEO
See MENDEZ 10	

$\Gamma(K^+\eta)/\Gamma(K^+K_S^0)$	$\Gamma_{85}/\Gamma_{31}$
Unseen decay modes of the $\eta$ are included.	
VALUE (units 10 <sup>-2</sup> )	EVTs
<b>11.8±2.2±0.6</b>	222 ± 41
MENDEZ	10 CLEO
e <sup>+</sup> e <sup>-</sup> at 4170 MeV	

$\Gamma(K^+\eta)/\Gamma(\eta\pi^+)$   $\Gamma_{85}/\Gamma_{68}$ 

VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.9 \pm 1.5 \pm 0.4$	$113 \pm 18$	ADAMS	07A	CLEO See MENDEZ 10

 $\Gamma(K^+\omega)/\Gamma_{\text{total}}$   $\Gamma_{86}/\Gamma$ Unseen decay modes of the  $\omega$  are included.

VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.24</b>	90	GE	09A	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(K^+\eta'(958))/\Gamma(K^+K_S^0)$   $\Gamma_{87}/\Gamma_{31}$ Unseen decay modes of the  $\eta'(958)$  are included.

VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>11.8 \pm 3.6 \pm 0.7</math></b>	$56 \pm 17$	MENDEZ	10	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(K^+\eta'(958))/\Gamma(\eta'(958)\pi^+)$   $\Gamma_{87}/\Gamma_{77}$ 

VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.2 \pm 1.3 \pm 0.3$	$28 \pm 9$	ADAMS	07A	CLEO See MENDEZ 10

 $\Gamma(K^+\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{88}/\Gamma$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.66 \pm 0.04</math> OUR FIT</b>			
<b><math>0.654 \pm 0.033 \pm 0.025</math></b>	ONYISI 13	CLEO	$e^+e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.69 \pm 0.05 \pm 0.03$	<sup>1</sup> ALEXANDER 08	CLEO	See ONYISI 13
<sup>1</sup> ALEXANDER 08 uses single- and double-tagged events in an overall fit.			

 $\Gamma(K^+\pi^+\pi^-)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{88}/\Gamma_{33}$ 

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.120 \pm 0.007</math> OUR FIT</b>				Error includes scale factor of 1.1.
<b><math>0.127 \pm 0.007 \pm 0.014</math></b>	$567 \pm 31$	LINK	04F	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV

 $\Gamma(K^+\rho^0)/\Gamma(K^+\pi^+\pi^-)$   $\Gamma_{89}/\Gamma_{88}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.3883 \pm 0.0531 \pm 0.0261</math></b>	LINK	04F	FOCS Dalitz fit, 567 evts

 $\Gamma(K^+\rho(1450)^0, \rho^0 \rightarrow \pi^+\pi^-)/\Gamma(K^+\pi^+\pi^-)$   $\Gamma_{90}/\Gamma_{88}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.1062 \pm 0.0351 \pm 0.0104</math></b>	LINK	04F	FOCS Dalitz fit, 567 evts

 $\Gamma(K^*(892)^0\pi^+, K^{*0} \rightarrow K^+\pi^-)/\Gamma(K^+\pi^+\pi^-)$   $\Gamma_{91}/\Gamma_{88}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.2164 \pm 0.0321 \pm 0.0114</math></b>	LINK	04F	FOCS Dalitz fit, 567 evts

 $\Gamma(K^*(1410)^0\pi^+, K^{*0} \rightarrow K^+\pi^-)/\Gamma(K^+\pi^+\pi^-)$   $\Gamma_{92}/\Gamma_{88}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.1882 \pm 0.0403 \pm 0.0122</math></b>	LINK	04F	FOCS Dalitz fit, 567 evts

 $\Gamma(K^*(1430)^0\pi^+, K^{*0} \rightarrow K^+\pi^-)/\Gamma(K^+\pi^+\pi^-)$   $\Gamma_{93}/\Gamma_{88}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0765 \pm 0.0500 \pm 0.0170</math></b>	LINK	04F	FOCS Dalitz fit, 567 evts

 $\Gamma(K^+\pi^+\pi^- \text{ nonresonant})/\Gamma(K^+\pi^+\pi^-)$   $\Gamma_{94}/\Gamma_{88}$ 

This is the “fit fraction” from the Dalitz-plot analysis.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.1588 \pm 0.0492 \pm 0.0153</math></b>	LINK	04F	FOCS Dalitz fit, 567 evts

 $\Gamma(K^0\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{95}/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>1.00 \pm 0.18 \pm 0.04</math></b>	$44 \pm 8$	NAIK	09A	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(K_S^0 2\pi^+\pi^-)/\Gamma(K_S^0 K^0 2\pi^+\pi^-)$   $\Gamma_{96}/\Gamma_{47}$ 

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.18 \pm 0.04 \pm 0.05</math></b>	$179 \pm 36$	LINK	08	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV

 $\Gamma(K^+\omega\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{97}/\Gamma$ Unseen decay modes of the  $\omega$  are included.

VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.82</b>	90	GE	09A	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(K^+\omega\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{98}/\Gamma$ Unseen decay modes of the  $\omega$  are included.

VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.54</b>	90	GE	09A	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(K^+\omega\eta)/\Gamma_{\text{total}}$   $\Gamma_{99}/\Gamma$ Unseen decay modes of the  $\omega$  and  $\eta$  are included.

VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.79</b>	90	GE	09A	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(2K^+K^-)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{100}/\Gamma_{33}$ 

VALUE (units $10^{-3}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>4.0 \pm 0.3 \pm 0.2</math></b>	$748 \pm 60$	DEL-AMO-SA..11g	BABR	$e^+e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.95 \pm 2.12 \pm 2.24$	31	LINK	02i	FOCS $\gamma A, \approx 180$ GeV

 $\Gamma(\phi K^+, \phi \rightarrow K^+K^-)/\Gamma(2K^+K^-)$   $\Gamma_{101}/\Gamma_{100}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.41 \pm 0.08 \pm 0.03</math></b>	DEL-AMO-SA..11g	BABR	$e^+e^- \approx \tau(4S)$

## Doubly Cabibbo-suppressed modes

 $\Gamma(2K^+\pi^-)/\Gamma(K^+K^-\pi^+)$   $\Gamma_{102}/\Gamma_{33}$ 

VALUE (units $10^{-3}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2.33 \pm 0.23</math> OUR AVERAGE</b>				
$2.3 \pm 0.3 \pm 0.2$	$356 \pm 52$	DEL-AMO-SA..11g	BABR	$e^+e^- \approx \tau(4S)$
$2.29 \pm 0.28 \pm 0.12$	$281 \pm 34$	KO	09	BELL $e^+e^-$ at $\tau(4S)$
$5.2 \pm 1.7 \pm 1.1$	$27 \pm 9$	LINK	05k	FOCS $<0.78\%$ , CL = 90%

 $\Gamma(K^+K^*(892)^0, K^{*0} \rightarrow K^+\pi^-)/\Gamma(2K^+\pi^-)$   $\Gamma_{103}/\Gamma_{102}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.47 \pm 0.22 \pm 0.15</math></b>	DEL-AMO-SA..11g	BABR	$e^+e^- \approx \tau(4S)$

## Baryon-antibaryon mode

 $\Gamma(p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{104}/\Gamma$ 

This is the only baryonic mode allowed kinematically.

VALUE (units $10^{-3}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>1.30 \pm 0.36 \pm 0.12</math></b>	$13.0 \pm 3.6$	ATHAR	08	CLEO $e^+e^-$ , $E_{\text{cm}} \approx 4170$ MeV

## Rare or forbidden modes

 $\Gamma(\pi^+e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{105}/\Gamma$ This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;13 \times 10^{-6}</math></b>	90	LEES	11g	BABR $e^+e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.2 \times 10^{-5}$	90	<sup>1</sup> RUBIN	10	CLEO $e^+e^-$ at 4170 MeV
$<27 \times 10^{-5}$	90	AITALA	99g	E791 $\pi^- N$ 500 GeV

<sup>1</sup>This RUBIN 10 limit is for the  $e^+e^-$  mass in the continuum away from the  $\phi(1020)$ . See the next data block. $\Gamma(\pi^+\phi, \phi \rightarrow e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{106}/\Gamma$ This is *not* a test for the  $\Delta C=1$  weak neutral current, but leads to the  $\pi^+e^+e^-$  final state.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>(6^{+8}_{-4} \pm 1) \times 10^{-6}</math></b>	3	RUBIN	10	CLEO $e^+e^-$ at 4170 MeV

 $\Gamma(\pi^+\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{107}/\Gamma$ This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;4.1 \times 10^{-7}</math></b>	90	AAIJ	13AF	LHCB $p\bar{p}$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4.3 \times 10^{-5}$	90	LEES	11g	BABR $e^+e^- \approx \tau(4S)$
$<2.6 \times 10^{-5}$	90	LINK	03F	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV
$<1.4 \times 10^{-4}$	90	AITALA	99g	E791 $\pi^- N$ 500 GeV
$<4.3 \times 10^{-4}$	90	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV

 $\Gamma(K^+e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{108}/\Gamma$ A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.7 \times 10^{-6}</math></b>	90	LEES	11g	BABR $e^+e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5.2 \times 10^{-5}$	90	RUBIN	10	CLEO $e^+e^-$ at 4170 MeV
$<1.6 \times 10^{-3}$	90	AITALA	99g	E791 $\pi^- N$ 500 GeV

 $\Gamma(K^+\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{109}/\Gamma$ A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;21 \times 10^{-6}</math></b>	90	LEES	11g	BABR $e^+e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<3.6 \times 10^{-5}$	90	LINK	03F	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV
$<1.4 \times 10^{-4}$	90	AITALA	99g	E791 $\pi^- N$ 500 GeV
$<5.9 \times 10^{-4}$	90	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV

 $\Gamma(K^*(892)^+\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{110}/\Gamma$ A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.4 \times 10^{-3}</math></b>	90	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV

# Meson Particle Listings

$D_s^\pm$

$\Gamma(\pi^+ e^+ \mu^-)/\Gamma_{\text{total}}$ A test of lepton-family-number conservation.					$\Gamma_{111}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<12 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	

$\Gamma(\pi^+ e^- \mu^+)/\Gamma_{\text{total}}$ A test of lepton-family-number conservation.					$\Gamma_{112}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<20 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	

$\Gamma(K^+ e^+ \mu^-)/\Gamma_{\text{total}}$ A test of lepton-family-number conservation.					$\Gamma_{113}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<14 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	

$\Gamma(K^+ e^- \mu^+)/\Gamma_{\text{total}}$ A test of lepton-family-number conservation.					$\Gamma_{114}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<9.7 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	

$\Gamma(\pi^- 2e^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{115}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<4.1 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.8 \times 10^{-5}$	90	RUBIN	10	CLEO $e^+ e^-$ at 4170 MeV	
$<69 \times 10^{-5}$	90	AITALA	99G	E791 $\pi^- N$ 500 GeV	

$\Gamma(\pi^- 2\mu^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{116}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.2 \times 10^{-7}$	90	AAIJ	13AF	LHCB $pp$ at 7 TeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.4 \times 10^{-5}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	
$<2.9 \times 10^{-5}$	90	LINK	03F	FOCS $\gamma A, \overline{E}_\gamma \approx 180$ GeV	
$<8.2 \times 10^{-5}$	90	AITALA	99G	E791 $\pi^- N$ 500 GeV	
$<4.3 \times 10^{-4}$	90	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV	

$\Gamma(\pi^- e^+ \mu^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{117}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<8.4 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<7.3 \times 10^{-4}$	90	AITALA	99G	E791 $\pi^- N$ 500 GeV	

$\Gamma(K^- 2e^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{118}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<5.2 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.7 \times 10^{-5}$	90	RUBIN	10	CLEO $e^+ e^-$ at 4170 MeV	
$<63 \times 10^{-5}$	90	AITALA	99G	E791 $\pi^- N$ 500 GeV	

$\Gamma(K^- 2\mu^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{119}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.3 \times 10^{-5}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	
$<1.3 \times 10^{-5}$	90	LINK	03F	FOCS $\gamma A, \overline{E}_\gamma \approx 180$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.8 \times 10^{-4}$	90	AITALA	99G	E791 $\pi^- N$ 500 GeV	
$<5.9 \times 10^{-4}$	90	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV	

$\Gamma(K^- e^+ \mu^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{120}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6.1 \times 10^{-6}$	90	LEES	11G	BABR $e^+ e^- \approx \mathcal{T}(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<6.8 \times 10^{-4}$	90	AITALA	99G	E791 $\pi^- N$ 500 GeV	

$\Gamma(K^*(892)-2\mu^+)/\Gamma_{\text{total}}$ A test of lepton-number conservation.					$\Gamma_{121}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.4 \times 10^{-3}$	90	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV	

## $D_s^+ - D_s^-$ CP-VIOLATING DECAY-RATE ASYMMETRIES

This is the difference between  $D_s^+$  and  $D_s^-$  partial widths for the decay to state  $f$ , divided by the sum of the widths:

$A_{CP}(f)=[\Gamma(D_s^+ \rightarrow f)-\Gamma(D_s^- \rightarrow \bar{f})]/[\Gamma(D_s^+ \rightarrow f)+\Gamma(D_s^- \rightarrow \bar{f})]$ .

$A_{CP}(\mu^\pm \nu)$ in $D_s^+ \rightarrow \mu^+ \nu, D_s^- \rightarrow \mu^- \bar{\nu}_\mu$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
$4.8 \pm 6.1$	ALEXANDER	09	CLEO $e^+ e^-$ at 4170 MeV	

$A_{CP}(K^\pm K_S^0)$ in $D_s^\pm \rightarrow K^\pm K_S^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.08 \pm 0.26</math> OUR AVERAGE</b>				
$-0.05 \pm 0.23 \pm 0.24$	288k	<sup>1</sup> LEES	13E	BABR $e^+ e^-$ at $\mathcal{T}(4S)$
$2.6 \pm 1.5 \pm 0.6$		ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV
$0.12 \pm 0.36 \pm 0.22$		KO	10	BELL $e^+ e^- \approx \mathcal{T}(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.7 \pm 1.8 \pm 0.9$	4.0k	MENDEZ	10	CLEO See ONYISI 13
$4.9 \pm 2.1 \pm 0.9$		ALEXANDER	08	CLEO See MENDEZ 10

<sup>1</sup>LEES 13E finds that after subtracting the contribution due to  $K^0 - \bar{K}^0$  mixing, the  $CP$  asymmetry is  $(+0.28 \pm 0.23 \pm 0.24)\%$ .

$A_{CP}(K^+ K^- \pi^\pm)$ in $D_s^\pm \rightarrow K^+ K^- \pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-0.5 \pm 0.8 \pm 0.4</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.3 \pm 1.1 \pm 0.8$	ALEXANDER	08	CLEO See ONYISI 13	

$A_{CP}(\phi \pi^\pm)$ in $D_s^\pm \rightarrow \phi \pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-0.38 \pm 0.26 \pm 0.08</math></b>	ABAZOV	14B	D0 $p\bar{p}$ at 1.96 TeV	

$A_{CP}(K^\pm K_S^0 \pi^0)$ in $D_s^\pm \rightarrow K^\pm K_S^0 \pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-1.6 \pm 6.0 \pm 1.1</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	

$A_{CP}(2K_S^0 \pi^\pm)$ in $D_s^\pm \rightarrow 2K_S^0 \pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>3.1 \pm 5.2 \pm 0.6</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	

$A_{CP}(K^+ K^- \pi^\pm \pi^0)$ in $D_s^\pm \rightarrow K^+ K^- \pi^\pm \pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>0.0 \pm 2.7 \pm 1.2</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-5.9 \pm 4.2 \pm 1.2$	ALEXANDER	08	CLEO See ONYISI 13	

$A_{CP}(K^\pm K_S^0 \pi^+ \pi^-)$ in $D_s^\pm \rightarrow K^\pm K_S^0 \pi^+ \pi^-$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-5.7 \pm 5.3 \pm 0.9</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	

$A_{CP}(K_S^0 K^\mp 2\pi^\pm)$ in $D_s^\pm \rightarrow K_S^0 K^\mp 2\pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>4.1 \pm 2.7 \pm 0.9</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.7 \pm 3.6 \pm 1.1$	ALEXANDER	08	CLEO See ONYISI 13	

$A_{CP}(\pi^+ \pi^- \pi^\pm)$ in $D_s^\pm \rightarrow \pi^+ \pi^- \pi^\pm$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-0.7 \pm 3.0 \pm 0.6</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.0 \pm 4.6 \pm 0.7$	ALEXANDER	08	CLEO See ONYISI 13	

$A_{CP}(\pi^\pm \eta)$ in $D_s^\pm \rightarrow \pi^\pm \eta$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 3.0 \pm 0.8</math></b>		ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-4.6 \pm 2.9 \pm 0.3$	2.5k	MENDEZ	10	CLEO See ONYISI 13
$-8.2 \pm 5.2 \pm 0.8$		ALEXANDER	08	CLEO See MENDEZ 10

$A_{CP}(\pi^\pm \eta')$ in $D_s^\pm \rightarrow \pi^\pm \eta'$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-2.2 \pm 2.2 \pm 0.6</math></b>		ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-6.1 \pm 3.0 \pm 0.3$	1.4k	MENDEZ	10	CLEO See ONYISI 13
$-5.5 \pm 3.7 \pm 1.2$		ALEXANDER	08	CLEO See MENDEZ 10

$A_{CP}(\eta \pi^\pm \pi^0)$ in $D_s^\pm \rightarrow \eta \pi^\pm \pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-0.5 \pm 3.9 \pm 2.0</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	

$A_{CP}(\eta' \pi^\pm \pi^0)$ in $D_s^\pm \rightarrow \eta' \pi^\pm \pi^0$				
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b><math>-0.4 \pm 7.4 \pm 1.9</math></b>	ONYISI	13	CLEO $e^+ e^-$ at 4.17 GeV	

$A_{CP}(K^\pm \pi^0)$ in $D_s^\pm \rightarrow K^\pm \pi^0$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-26.6 \pm 23.8 \pm 0.9</math></b>	$202 \pm 70$	MENDEZ	10	CLEO $e^+ e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2 \pm 29$		ADAMS	07A	CLEO See MENDEZ 10

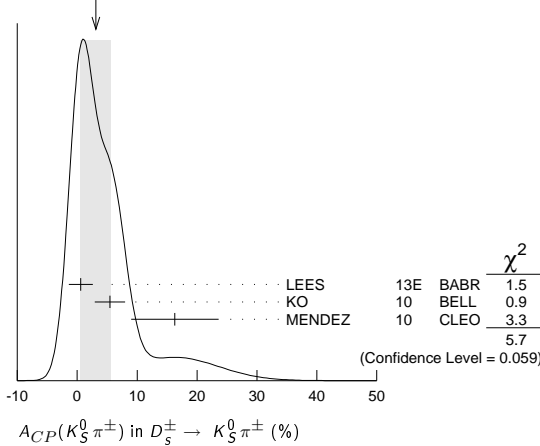


$A_{CP}(\bar{K}^0/K^0\pi^\pm)$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.4 ± 0.5 OUR AVERAGE</b>				
0.38 ± 0.46 ± 0.17	121k	<sup>1</sup> AAIJ	14BD LHCb	$pp$ at 7, 8 TeV
0.3 ± 2.0 ± 0.3	14k	LEES	13E BABR	$e^+e^-$ at $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.61 ± 0.83 ± 0.14	26k	AAIJ	13W LHCb	See AAIJ 14BD
<sup>1</sup> AAIJ 14BD reports its result as $A_{CP}(D_s^\pm \rightarrow K_S^0 K^\pm)$ with $CP$ -violation effects in the $K^0 - \bar{K}^0$ system subtracted. It also measures $A_{CP}(D^\pm \rightarrow \bar{K}^0/K^0 K^\pm) + A_{CP}(D_s^\pm \rightarrow \bar{K}^0/K^0 \pi^\pm) = (0.41 \pm 0.49 \pm 0.26)\%$ .				

 $A_{CP}(K_S^0\pi^\pm)$  in  $D_s^\pm \rightarrow K_S^0\pi^\pm$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.1 ± 2.6 OUR AVERAGE</b>				Error includes scale factor of 1.7. See the ideogram below.
0.6 ± 2.0 ± 0.3	14k	LEES	13E BABR	$e^+e^-$ at $\Upsilon(4S)$
5.45 ± 2.50 ± 0.33		KO	10 BELL	$e^+e^- \approx \Upsilon(4S)$
16.3 ± 7.3 ± 0.3	0.4k	MENDEZ	10 CLEO	$e^+e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
27 ± 11		ADAMS	07A CLEO	See MENDEZ 10

WEIGHTED AVERAGE  
3.1 ± 2.6 (Error scaled by 1.7) $A_{CP}(K^\pm\pi^+\pi^-)$  in  $D_s^\pm \rightarrow K^\pm\pi^+\pi^-$ 

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>4.5 ± 4.8 ± 0.6</b>	ONYISI	13 CLEO	$e^+e^-$ at 4.17 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11.2 ± 7.0 ± 0.9	ALEXANDER	08 CLEO	See ONYISI 13

 $A_{CP}(K^\pm\eta)$  in  $D_s^\pm \rightarrow K^\pm\eta$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.3 ± 15.2 ± 0.9</b>	222 ± 41	MENDEZ	10 CLEO	$e^+e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-20 ± 18		ADAMS	07A CLEO	See MENDEZ 10

 $A_{CP}(K^\pm\eta(958))$  in  $D_s^\pm \rightarrow K^\pm\eta(958)$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.0 ± 18.9 ± 0.9</b>	56 ± 17	MENDEZ	10 CLEO	$e^+e^-$ at 4170 MeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-17 ± 37		ADAMS	07A CLEO	See MENDEZ 10

## CP VIOLATING ASYMMETRIES OF P-ODD (T-ODD) MOMENTS

 $A_{T\text{viol}}(K_S^0 K^\pm\pi^+\pi^-)$  in  $D_s^\pm \rightarrow K_S^0 K^\pm\pi^+\pi^-$ 

$C_T \equiv \bar{p}_{K^+} \cdot (\bar{p}_{\pi^+} \times \bar{p}_{\pi^-})$  is a parity-odd correlation of the  $K^+$ ,  $\pi^+$ , and  $\pi^-$  momenta for the  $D_s^\pm$ .  $\bar{C}_T \equiv \bar{p}_{K^-} \cdot (\bar{p}_{\pi^-} \times \bar{p}_{\pi^+})$  is the corresponding quantity for the  $D_s^-$ . Then

$$A_T \equiv [\Gamma(C_T > 0) - \Gamma(C_T < 0)] / [\Gamma(C_T > 0) + \Gamma(C_T < 0)], \text{ and}$$

$$\bar{A}_T \equiv [\Gamma(-\bar{C}_T > 0) - \Gamma(-\bar{C}_T < 0)] / [\Gamma(-\bar{C}_T > 0) + \Gamma(-\bar{C}_T < 0)], \text{ and}$$

$A_{T\text{viol}} \equiv \frac{1}{2}(A_T - \bar{A}_T)$ .  $C_T$  and  $\bar{C}_T$  are commonly referred to as  $T$ -odd moments, because they are odd under  $T$  reversal. However, the  $T$ -conjugate process  $K_S^0 K^\pm\pi^+\pi^- \rightarrow D_s^\pm$  is not accessible, while the  $P$ -conjugate process is.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-13.6 ± 7.7 ± 3.4</b>	29.8 ± 0.3k	LEES	11E BABR	$e^+e^- \approx \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-36 ± 67 ± 23	508 ± 34	LINK	05E FOCS	$\gamma A, \bar{E}_\gamma \approx 180 \text{ GeV}$

 $D_s^\pm \rightarrow \phi\ell^+\nu_\ell$  FORM FACTORS $r_2 \equiv A_2(0)/A_1(0)$  in  $D_s^\pm \rightarrow \phi\ell^+\nu_\ell$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.84 ± 0.11 OUR AVERAGE</b>				Error includes scale factor of 2.4.
0.816 ± 0.036 ± 0.030	25 ± 0.5k	<sup>1</sup> AUBERT	08AN BABR	$\phi e^+ \nu_e$
0.713 ± 0.202 ± 0.284	793	LINK	04C FOCS	$\phi\mu^+ \nu_\mu$
1.57 ± 0.25 ± 0.19	271	AITALA	99D E791	$\phi e^+ \nu_e, \phi\mu^+ \nu_\mu$
1.4 ± 0.5 ± 0.3	308	VERY	94B CLE2	$\phi e^+ \nu_e$
1.1 ± 0.8 ± 0.1	90	FRABETTI	94F E687	$\phi\mu^+ \nu_\mu$
2.1 $\begin{smallmatrix} +0.6 \\ -0.5 \end{smallmatrix}$ ± 0.2	19	KODAMA	93 E653	$\phi\mu^+ \nu_\mu$

<sup>1</sup> To compare with previous measurements, this AUBERT 08AN value is from a fit that fixes the pole masses at  $m_A = 2.5 \text{ GeV}/c^2$  and  $m_V = 2.1 \text{ GeV}/c^2$ . A simultaneous fit to  $r_2$ ,  $r_V$ ,  $r_0$  (a significant  $s$ -wave contribution) and  $m_A$ , gives  $r_2 = 0.763 \pm 0.071 \pm 0.065$ .

 $r_V \equiv V(0)/A_1(0)$  in  $D_s^\pm \rightarrow \phi\ell^+\nu_\ell$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.80 ± 0.08 OUR AVERAGE</b>				
1.807 ± 0.046 ± 0.065	25 ± 0.5k	<sup>1</sup> AUBERT	08AN BABR	$\phi e^+ \nu_e$
1.549 ± 0.250 ± 0.148	793	LINK	04C FOCS	$\phi\mu^+ \nu_\mu$
2.27 ± 0.35 ± 0.22	271	AITALA	99D E791	$\phi e^+ \nu_e, \phi\mu^+ \nu_\mu$
0.9 ± 0.6 ± 0.3	308	VERY	94B CLE2	$\phi e^+ \nu_e$
1.8 ± 0.9 ± 0.2	90	FRABETTI	94F E687	$\phi\mu^+ \nu_\mu$
2.3 $\begin{smallmatrix} +1.1 \\ -0.9 \end{smallmatrix}$ ± 0.4	19	KODAMA	93 E653	$\phi\mu^+ \nu_\mu$

<sup>1</sup> To compare with previous measurements, this AUBERT 08AN value is from a fit that fixes the pole masses at  $m_A = 2.5 \text{ GeV}/c^2$  and  $m_V = 2.1 \text{ GeV}/c^2$ . A simultaneous fit to  $r_2$ ,  $r_V$ ,  $r_0$  (a significant  $s$ -wave contribution) and  $m_A$ , gives  $r_V = 1.849 \pm 0.060 \pm 0.055$ .

 $\Gamma_L/\Gamma_T$  in  $D_s^\pm \rightarrow \phi\ell^+\nu_\ell$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.72 ± 0.18 OUR AVERAGE</b>				
1.0 ± 0.3 ± 0.2	308	VERY	94B CLE2	$\phi e^+ \nu_e$
1.0 ± 0.5 ± 0.1	90	<sup>1</sup> FRABETTI	94F E687	$\phi\mu^+ \nu_\mu$
0.54 ± 0.21 ± 0.10	19	<sup>1</sup> KODAMA	93 E653	$\phi\mu^+ \nu_\mu$

<sup>1</sup> FRABETTI 94F and KODAMA 93 evaluate  $\Gamma_L/\Gamma_T$  for a lepton mass of zero.

 $D_s^\pm$  REFERENCES

ABLIKIM	15Z	PL B750 466	M. Ablikim <i>et al.</i>	(BES III Collab.)
HIETALA	15	PR D92 012009	J. Hietala <i>et al.</i>	(MINN, LUTH, OXF)
LEES	15D	PR D91 019901 (err.)	J.P. Lees <i>et al.</i>	(BABAR Collab.)
AAIJ	14BD	JHEP 1410 025	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	14B	PRL 112 111804	V.M. Abazov <i>et al.</i>	(DO Collab.)
AAIJ	13AF	PL B724 203	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13V	JHEP 1306 065	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13W	JHEP 1306 112	R. Aaij <i>et al.</i>	(LHCb Collab.)
LEES	13E	PR D87 052012	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ONYISI	13	PR D88 032009	P.U.E. Onyisi <i>et al.</i>	(CLEO Collab.)
ZUPANC	13	JHEP 1309 139	A. Zupanc <i>et al.</i>	(BELLE Collab.)
DEL-AMO-SA...	11G	PR D83 052001	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
LEES	11E	PR D84 031103	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	11G	PR D84 072006	J.P. Lees <i>et al.</i>	(BABAR Collab.)
MARTIN	11	PR D84 012005	L. Martin <i>et al.</i>	(CLEO Collab.)
ASNER	10	PR D81 052007	D.M. Asner <i>et al.</i>	(CLEO Collab.)
DEL-AMO-SA...	10J	PR D82 091103	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
Also		PR D91 019901 (err.)	J.P. Lees <i>et al.</i>	(BABAR Collab.)
KO	10	PRL 104 181602	B.R. Ko <i>et al.</i>	(BELLE Collab.)
MENDEZ	10	PR D81 052013	H. Mendez <i>et al.</i>	(CLEO Collab.)
RUBIN	10	PR D82 092007	P. Rubin <i>et al.</i>	(CLEO Collab.)
ALEXANDER	09	PR D79 052001	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
AUBERT	09O	PR D79 032003	B. Aubert <i>et al.</i>	(BABAR Collab.)
DOBBS	09	PR D79 122008	S. Dobbs <i>et al.</i>	(CLEO Collab.)
ECKLUND	09	PR D80 052009	K.M. Ecklund <i>et al.</i>	(CLEO Collab.)
GE	09A	PR D80 051102	J.Y. Ge <i>et al.</i>	(CLEO Collab.)
KO	09	PRL 102 221802	B.R. Ko <i>et al.</i>	(BELLE Collab.)
MITCHELL	09A	PR D79 072008	R.E. Mitchell <i>et al.</i>	(CLEO Collab.)
NAIK	09A	PR D80 112004	P. Naik <i>et al.</i>	(CLEO Collab.)
ONYISI	09	PR D79 052002	P.U.E. Onyisi <i>et al.</i>	(CLEO Collab.)
WON	09	PR D80 111101	E. Won <i>et al.</i>	(BELLE Collab.)
YELTON	09	PR D80 052007	J. Yelton <i>et al.</i>	(CLEO Collab.)
ALEXANDER	08	PRL 100 161804	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
ATHAR	08	PRL 100 181802	S.B. Athar <i>et al.</i>	(CLEO Collab.)
AUBERT	08AN	PR D78 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
ECKLUND	08	PRL 100 161801	K.M. Ecklund <i>et al.</i>	(CLEO Collab.)
KLEMP	08	EPJ C55 39	E. Klemp, M. Matveev, A.V. Sarantsev	(BONN+)
LINK	08	PL B660 147	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
WIDHALM	08	PRL 100 241801	L. Widhalm <i>et al.</i>	(BELLE Collab.)
ADAMS	07A	PRL 99 191805	G.S. Adams <i>et al.</i>	(CLEO Collab.)
AUBERT	07V	PRL 98 141801	B. Aubert <i>et al.</i>	(BABAR Collab.)
PEDLAR	07A	PR D76 072002	T.K. Pedlar <i>et al.</i>	(CLEO Collab.)
Also		PRL 99 071802	M. Artuso <i>et al.</i>	(CLEO Collab.)
AUBERT	06N	PR D74 031103	B. Aubert <i>et al.</i>	(BABAR Collab.)
HUANG	06B	PR D74 112005	G.S. Huang <i>et al.</i>	(CLEO Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
AUBERT	05V	PR D71 091104	B. Aubert <i>et al.</i>	(BABAR Collab.)
LINK	05E	PL B622 239	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	05J	PRL 95 052003	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	05K	PL B624 166	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	04	PL B585 200	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	04C	PL B586 183	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	04D	PL B586 191	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	04F	PL B601 10	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ACOSTA	03D	PR D68 072004	D. Acosta <i>et al.</i>	(FNAL CDF-II Collab.)
ANISOVICH	03	EPJ A16 229	V.V. Anisovich <i>et al.</i>	
LINK	03D	PL B561 225	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	03F	PL B572 21	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
AUBERT	02G	PR D65 091104	B. Aubert <i>et al.</i>	(BABAR Collab.)
HEISTER	02I	PL B528 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
LINK	02I	PL B541 227	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)

# Meson Particle Listings

## $D_s^\pm, D_s^{*\pm}$

LINK	02J	PL B541 243	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ABBIENDI	01L	PL B516 236	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
AITALA	01A	PRL 86 765	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
IORI	01	PL B523 22	M. Iori <i>et al.</i>	(FNAL SELEX Collab.)
LINK	01C	PRL 87 162001	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
ALEXANDROV	00	PL B478 31	Y. Alexandrov <i>et al.</i>	(CERN BEATRICE Collab.)
AITALA	99	PL B445 449	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	99D	PL B450 294	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
AITALA	99G	PL B462 401	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
BONVICINI	99	PRL 82 4566	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
CHADHA	98	PR D58 032002	M. Chada <i>et al.</i>	(CLEO Collab.)
JESSOP	98	PR D58 052002	C.P. Jessop <i>et al.</i>	(CLEO Collab.)
ACCIARRI	97F	PL B396 327	M. Acciari <i>et al.</i>	(L3 Collab.)
BALEST	97	PRL 79 1436	R. Balest <i>et al.</i>	(CLEO Collab.)
FRABETTI	97C	PL B401 131	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	97D	PL B407 79	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ARTUSO	96	PL B378 364	M. Artuso <i>et al.</i>	(CLEO Collab.)
BAI	95C	PR D52 3781	J.Z. Bai <i>et al.</i>	(BES Collab.)
BRANDENB...	95	PRL 75 3804	G.W. Brandenburg <i>et al.</i>	(CLEO Collab.)
FRABETTI	95B	PL B351 591	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	95	PL B345 85	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ACOSTA	94	PR D49 5690	D. Acosta <i>et al.</i>	(CLEO Collab.)
AVERY	94B	PL B337 405	P. Avery <i>et al.</i>	(CLEO Collab.)
BROWN	94	PR D50 1884	D. Brown <i>et al.</i>	(CLEO Collab.)
BUTLER	94	PL B324 255	F. Butler <i>et al.</i>	(CLEO Collab.)
FRABETTI	94F	PL B328 187	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	93F	PRL 71 827	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	93G	PL B313 253	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	93	PL B309 483	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ALBRECHT	92B	ZPHY C53 361	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	92	PRL 68 1275	J. Alexander <i>et al.</i>	(CLEO Collab.)
AVERY	92	PRL 68 1279	P. Avery <i>et al.</i>	(CLEO Collab.)
BARLAG	92C	ZPHY C55 383	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
Also		ZPHY C48 29	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
FRABETTI	92	PL B281 167	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALBRECHT	91	PL B255 634	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALVAREZ	91	PL B255 639	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
ALBRECHT	90D	PL B245 315	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	90B	PRL 65 1531	J. Alexander <i>et al.</i>	(CLEO Collab.)
BARLAG	90C	ZPHY C46 563	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
FRABETTI	90	PL B251 639	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ANJOS	89E	PL B223 267	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
CHEN	89	PL B226 192	W.Y. Chen <i>et al.</i>	(CLEO Collab.)
ALBRECHT	88	PL B207 349	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	88	PRL 60 897	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
RAAB	88	PR D37 2391	J.R. Raab <i>et al.</i>	(FNAL E691 Collab.)
BECKER	87B	PL B184 277	H. Becker <i>et al.</i>	(NA11 and NA32 Collabs.)
BLAYLOCK	87	PRL 58 2171	G.T. Blaylock <i>et al.</i>	(Mark III Collab.)
USHIDA	86	PRL 56 1767	N. Ushida <i>et al.</i>	(FNAL E531 Collab.)
ALBRECHT	85D	PL 153B 343	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DERRICK	85B	PRL 54 2568	M. Derrick <i>et al.</i>	(HRS Collab.)
AIHARA	84D	PRL 53 2465	H. Aihara <i>et al.</i>	(TPC Collab.)
ALTHOFF	84	PL 136B 130	M. Althoff <i>et al.</i>	(TASSO Collab.)
BAILEY	84	PL 139B 320	R. Bailey <i>et al.</i>	(ACCMOR Collab.)
CHEN	83C	PRL 51 634	A. Chen <i>et al.</i>	(CLEO Collab.)

### OTHER RELATED PAPERS

RICHMAN	95	RMP 67 893	J.D. Richman, P.R. Burchat	(UCSB, STAN)
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$D_s^{*\pm}$

$$I(J^P) = 0(?^?)$$

$J^P$  is natural, width and decay modes consistent with  $1^-$ .

### $D_s^{*\pm}$ MASS

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2112.1±0.4 OUR FIT</b>			
<b>2106.6±2.1±2.7</b>	1	BLAYLOCK 87	MRK3 $e^+e^- \rightarrow D_s^\pm \gamma X$

1 Assuming  $D_s^\pm$  mass = 1968.7 ± 0.9 MeV.

$$m_{D_s^{*\pm}} - m_{D_s^\pm}$$

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>143.8 ± 0.4 OUR FIT</b>				
<b>143.9 ± 0.4 OUR AVERAGE</b>				
143.76± 0.39±0.40		GRONBERG 95	CLE2	$e^+e^-$
144.22± 0.47±0.37		BROWN 94	CLE2	$e^+e^-$
142.5 ± 0.8 ±1.5		2	ALBRECHT 88	ARG $e^+e^- \rightarrow D_s^\pm \gamma X$
139.5 ± 8.3 ±9.7	60	AIHARA 84D	TPC	$e^+e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
143.0 ±18.0	8	ASRATYAN 85	HLBC	FNAL 15-ft, $\nu^-2$ H
110 ±46		BRA NDELIK 79	DASP	$e^+e^- \rightarrow D_s^\pm \gamma X$

2 Result includes data of ALBRECHT 84B.

### $D_s^{*\pm}$ WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.9</b>	90	GRONBERG 95	CLE2	$e^+e^-$
< 4.5	90	ALBRECHT 88	ARG	$E_{\rm cm}^{\rm e^+e^-} = 10.2$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				

< 4.9	90	BROWN 94	CLE2	$e^+e^-$
<22	90	BLAYLOCK 87	MRK3	$e^+e^- \rightarrow D_s^\pm \gamma X$

### $D_s^{*+}$ DECAY MODES

$D_s^{*-}$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad D_s^+ \gamma$	(93.5±0.7) %
$\Gamma_2 \quad D_s^+ \pi^0$	( 5.8±0.7) %
$\Gamma_3 \quad D_s^+ e^+ e^-$	( 6.7±1.6) × 10 <sup>−3</sup>

### CONSTRAINED FIT INFORMATION

An overall fit to 2 branching ratios uses 3 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 0.0$  for 1 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_j \equiv \Gamma_j/\Gamma_{\rm total}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−97	
$x_3$	−19	−4
	$x_1$	$x_2$

### $D_s^{*+}$ BRANCHING RATIOS

$\Gamma(D_s^{*+} \gamma)/\Gamma_{\rm total}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>0.935±0.007 OUR FIT</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen		ASRATYAN 91	HLBC	$\overline{\nu}_\mu$ Ne	
seen		ALBRECHT 88	ARG	$e^+e^- \rightarrow D_s^\pm \gamma X$	
seen		AIHARA 84D			
seen		ALBRECHT 84B			
seen		BRANDELIK 79			

$\Gamma(D_s^{*+} \pi^0)/\Gamma(D_s^{*+} \gamma)$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_1$
<b>0.062±0.008 OUR FIT</b>					
<b>0.062±0.005 OUR AVERAGE</b>					
0.062±0.008±0.006		AUBERT,BE 05G	BABR	10.6 $e^+e^- \rightarrow$ hadrons	
0.062+0.020−0.018±0.022		GRONBERG 95	CLE2	$e^+e^-$	

$\Gamma(D_s^{*+} e^+ e^-)/\Gamma(D_s^{*+} \gamma)$	VALUE (units 10 <sup>−3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma_1$
<b>7.2±1.7 OUR FIT</b>						
<b>7.2+1.5−1.3±1.0</b>		38	CRONIN-HEN..12	CLEO	4.17 $e^+e^- \rightarrow$ hadrons	

### $D_s^{*+}$ REFERENCES

CRONIN-HEN...12	PR D86 072005	D. Cronin-Hennessey <i>et al.</i>	(CLEO Collab.)
AUBERT,BE 05G	PR D72 091101	B. Aubert <i>et al.</i>	(BABAR Collab.)
GRONBERG 95	PRL 75 3232	J. Gronberg <i>et al.</i>	(CLEO Collab.)
BROWN 94	PR D50 1884	D. Brown <i>et al.</i>	(CLEO Collab.)
ASRATYAN 91	PL B257 525	A.E. Asratyan <i>et al.</i>	(ITEP, BELG, SACL+)
ALBRECHT 88	PL B207 349	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BLAYLOCK 87	PRL 58 2171	G.T. Blaylock <i>et al.</i>	(Mark III Collab.)
ASRATYAN 85	PL 156B 441	A.E. Asratyan <i>et al.</i>	(ITEP, SERP)
AIHARA 84D	PRL 53 2465	H. Aihara <i>et al.</i>	(TPC Collab.)
ALBRECHT 84B	PL 146B 111	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BRANDELIK 79	PL 80B 412	R. Brandelik <i>et al.</i>	(DASP Collab.)

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# Meson Particle Listings

$$D_{s0}^*(2317)^\pm, D_{s1}(2460)^\pm$$

$D_{s0}^*(2317)^\pm$

$I(J^P) = 0(0^+)$   
 $J, P$  need confirmation.

AUBERT 06P and CHOI 15A do not observe neutral and doubly charged partners of the  $D_{s0}^*(2317)^\pm$ .

## $D_{s0}^*(2317)^\pm$ MASS

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2317.7±0.6 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>2318.0±1.0 OUR AVERAGE</b>	Error includes scale factor of 1.4.			
2319.6±0.2±1.4	3180	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup> → $D_s^+\pi^0 X$
2317.3±0.4±0.8	1022	<sup>1</sup> AUBERT	04E BABR	10.6 e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2317.2±1.3	88	<sup>2</sup> AUBERT,B	04s BABR	$B \rightarrow D_{s0}^{(*)}(2317)+\overline{D}^{(*)}$
2317.2±0.5±0.9	761	<sup>3</sup> MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
2316.8±0.4±3.0	1267±53	<sup>3,4</sup> AUBERT	03G BABR	10.6 e <sup>+</sup> e <sup>-</sup>
2317.6±1.3	273±33	<sup>3,5</sup> AUBERT	03G BABR	10.6 e <sup>+</sup> e <sup>-</sup>
2319.8±2.1±2.0	24	<sup>3</sup> KROKOVNY	03B BELL	10.6 e <sup>+</sup> e <sup>-</sup>
<sup>1</sup> Supersedes AUBERT 03G.				
<sup>2</sup> Systematic errors not evaluated.				
<sup>3</sup> Not independent of the corresponding $m_{D_{s0}^*(2317)} - m_{D_s}$ .				
<sup>4</sup> From $D_s^+ \rightarrow K^+ K^- \pi^+$ decay.				
<sup>5</sup> From $D_s^+ \rightarrow K^+ K^- \pi^+ \pi^0$ decay.				

## $m_{D_{s0}^*(2317)^\pm} - m_{D_s^\pm}$

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>349.4±0.6 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>349.2±0.7 OUR AVERAGE</b>				
348.7±0.5±0.7	761	MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
350.0±1.2±1.0	135	BESSION	03 CLE2	10.6 e <sup>+</sup> e <sup>-</sup>
351.3±2.1±1.9	24	<sup>6</sup> KROKOVNY	03B BELL	10.6 e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
349.6±0.4±3.0	1267	<sup>7,8</sup> AUBERT	03G BABR	10.6 e <sup>+</sup> e <sup>-</sup>
350.2±1.3	273	<sup>9,10</sup> AUBERT	03G BABR	10.6 e <sup>+</sup> e <sup>-</sup>
<sup>6</sup> Recalculated by us using $m_{D_s^+} = 1968.5 \pm 0.6$ MeV.				
<sup>7</sup> From $D_s^+ \rightarrow K^+ K^- \pi^+$ decay.				
<sup>8</sup> Recalculated by us using $m_{D_s^+} = 1967.20 \pm 0.03$ MeV.				
<sup>9</sup> From $D_s^+ \rightarrow K^+ K^- \pi^+ \pi^0$ decay.				
<sup>10</sup> Recalculated by us using $m_{D_s^+} = 1967.4 \pm 0.2$ MeV. Systematic errors not estimated.				

## $D_{s0}^*(2317)^\pm$ WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.8</b>	95	3180	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup> → $D_s^+\pi^0 X$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 4.6	90	761	MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
<10			AUBERT	03G BABR	10.6 e <sup>+</sup> e <sup>-</sup>
< 7	90	135	BESSION	03 CLE2	10.6 e <sup>+</sup> e <sup>-</sup>

## $D_{s0}^*(2317)^\pm$ DECAY MODES

$D_{s0}^*(2317)^-$  modes are charge conjugates of modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D_s^+\pi^0$	seen
$\Gamma_2$ $D_s^+\gamma$	
$\Gamma_3$ $D_s^*(2112)^+\gamma$	
$\Gamma_4$ $D_s^+\gamma\gamma$	
$\Gamma_5$ $D_s^*(2112)^+\pi^0$	
$\Gamma_6$ $D_s^+\pi^+\pi^-$	
$\Gamma_7$ $D_s^+\pi^0\pi^0$	not seen

## $D_{s0}^*(2317)^\pm$ BRANCHING RATIOS

$\Gamma(D_s^+ \pi^0)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	$1540 \pm 62$	AUBERT	03G BABR	$10.6 e^+ e^-$

$\Gamma(D_s^+\gamma)/\Gamma(D_s^+\pi^0)$	$\Gamma_2/\Gamma_1$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.05</b>	90	MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.14	95	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup>
<0.052	90	BESSION	03 CLE2	10.6 e <sup>+</sup> e <sup>-</sup>

$\Gamma(D_s^*(2112)^+\gamma)/\Gamma(D_s^+\pi^0)$					$\Gamma_3/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.059	90	BESSION	03 CLE2	10.6 e <sup>+</sup> e <sup>-</sup>	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.16	95	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup>	
<0.18	90	MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>	

$\Gamma(D_s^+\gamma\gamma)/\Gamma(D_s^+\pi^0)$					$\Gamma_4/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.18</b>	95	AUBERT	06P	BABR	10.6 e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •					
not seen		AUBERT	03G	BABR	10.6 e <sup>+</sup> e <sup>-</sup>

$\Gamma(D_s^*(2112)^+\pi^0)/\Gamma(D_s^+\pi^0)$					$\Gamma_5/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.11</b>	90	BESSION	03	CLE2	10.6 e <sup>+</sup> e <sup>-</sup>

$\Gamma(D_s^+\pi^+\pi^-)/\Gamma(D_s^+\pi^0)$					$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.004	90	MIKAMI	04	BELL	10.6 e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.005	95	AUBERT	06P	BABR	10.6 e <sup>+</sup> e <sup>-</sup>
<0.019	90	BESSION	03	CLE2	10.6 e <sup>+</sup> e <sup>-</sup>

$\Gamma(D_s^+\pi^0\pi^0)/\Gamma(D_s^+\pi^0)$					$\Gamma_7/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.25</b>	95	AUBERT	06P	BABR	10.6 e <sup>+</sup> e <sup>-</sup>

## $D_{s0}^*(2317)^\pm$ REFERENCES

CHOI	15A	PR D91 092011	S.-K. Choi <i>et al.</i>	(BELLE Collab.)
AUBERT	06P	PR D74 032007	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	04E	PR D69 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	04S	PRL 93 181801	B. Aubert <i>et al.</i>	(BABAR Collab.)
MIKAMI	04	PRL 92 012002	Y. Mikami <i>et al.</i>	(BELLE Collab.)
AUBERT	03G	PRL 90 242001	B. Aubert <i>et al.</i>	(BABAR Collab.)
BESSION	03	PR D68 032002	D. Besson <i>et al.</i>	(CLEO Collab.)
KROKOVNY	03B	PRL 91 262002	P. Krokovny <i>et al.</i>	(BELLE Collab.)

$$D_{s1}(2460)^\pm$$

$$I(J^P) = 0(1^+)$$

## $D_{s1}(2460)^\pm$ MASS

The fit includes  $D^\pm, D^0, D_s^\pm, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^*(2460)^0$ , and  $D_{s1}(2536)^\pm$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2459.5±0.6 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>2459.6±0.9 OUR AVERAGE</b>	Error includes scale factor of 1.3.			
2460.1±0.2±0.8		<sup>1</sup> AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup>
2458.0±1.0±1.0	195	AUBERT	04E BABR	10.6 e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2459.5±1.2±3.7	920	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup> → $D_s^+\gamma X$
2458.6±1.0±2.5	560	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup> → $D_s^+\pi^0\gamma X$
2460.2±0.2±0.8	123	AUBERT	06P BABR	10.6 e <sup>+</sup> e <sup>-</sup> → $D_s^+\pi^+\pi^- X$
2458.9±1.5	112	<sup>2</sup> AUBERT,B	04s BABR	$B \rightarrow D_{s1}(2460)+\overline{D}^{(*)}$
2461.1±1.6	139	<sup>3</sup> AUBERT,B	04s BABR	$B \rightarrow D_{s1}(2460)+\overline{D}^{(*)}$
2456.5±1.3±1.3	126	<sup>4,5</sup> MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
2459.5±1.3±2.0	152	<sup>6,7</sup> MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
2459.9±0.9±1.6	60	<sup>6,7</sup> MIKAMI	04 BELL	10.6 e <sup>+</sup> e <sup>-</sup>
2459.2±1.6±2.0	57	KROKOVNY	03B BELL	10.6 e <sup>+</sup> e <sup>-</sup>

<sup>1</sup> The average of the values obtained from the  $D_s^+\gamma, D_s^+\pi^0\gamma, D_s^+\pi^+\pi^-$  final state.

<sup>2</sup> Systematic errors not evaluated. From the decay to  $D_s^{*+}\pi^0$ .

<sup>3</sup> Systematic errors not evaluated. From the decay to  $D_s^{*+}\gamma$ .

<sup>4</sup> Not independent of the corresponding  $m_{D_{s1}(2460)^\pm} - m_{D_s^\pm}$ .

<sup>5</sup> Using  $m_{D_s^{*+}} = 2112.4 \pm 0.7$  MeV.

<sup>6</sup> Not independent of the corresponding  $m_{D_{s1}(2460)^\pm} - m_{D_s^\pm}$ .

<sup>7</sup> Using  $m_{D_s^+} = 1968.5 \pm 0.6$  MeV.

# Meson Particle Listings

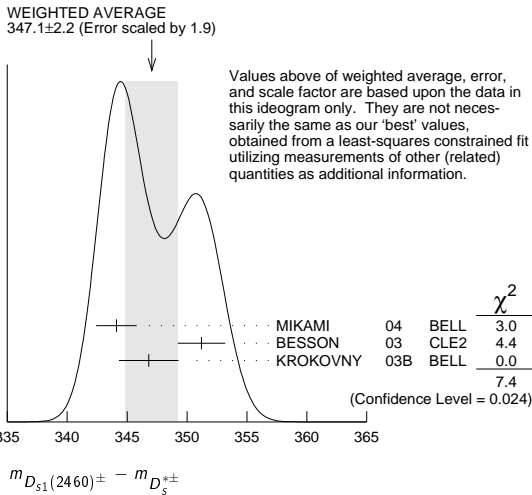
## $D_{s1}(2460)^{\pm}$

$$m_{D_{s1}(2460)^{\pm}} - m_{D_s^{\pm}}$$

The fit includes  $D^{\pm}, D^0, D_s^{\pm}, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^{*}(2460)^0$ , and  $D_{s1}(2536)^{\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>347.3±0.7 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>347.1±2.2 OUR AVERAGE</b>	Error includes scale factor of 1.9. See the ideogram below.			
344.1±1.3±1.1	126	MIKAMI	04 BELL	10.6 $e^+e^-$
351.2±1.7±1.0	41	BESSON	03 CLE2	10.6 $e^+e^-$
346.8±1.6±1.9	57	<sup>8</sup> KROKOVNY	03B BELL	10.6 $e^+e^-$

<sup>8</sup> Recalculated by us using  $m_{D_s^{*+}} = 2112.4 \pm 0.7$  MeV.



$$m_{D_{s1}(2460)^{\pm}} - m_{D_s^{\pm}}$$

The fit includes  $D^{\pm}, D^0, D_s^{\pm}, D^{*\pm}, D^{*0}, D_s^{*\pm}, D_1(2420)^0, D_2^{*}(2460)^0$ , and  $D_{s1}(2536)^{\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>491.2±0.6 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>491.3±1.4 OUR AVERAGE</b>				
491.0±1.3±1.9	152	<sup>9</sup> MIKAMI	04 BELL	10.6 $e^+e^-$
491.4±0.9±1.5	60	<sup>10</sup> MIKAMI	04 BELL	10.6 $e^+e^-$

<sup>9</sup> From the decay to  $D_s^{\pm}\gamma$ .  
<sup>10</sup> From the decay to  $D_s^{\pm}\pi^+\pi^-$ .

### $D_{s1}(2460)^{\pm}$ WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.5</b>	95	123	AUBERT	06P BABR	10.6 $e^+e^- \rightarrow D_s^+\pi^+\pi^-X$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 6.3	95	560	AUBERT	06P BABR	10.6 $e^+e^- \rightarrow D_s^+\pi^0\gamma X$
<10		195	AUBERT	04E BABR	10.6 $e^+e^-$
< 5.5	90	126	MIKAMI	04 BELL	10.6 $e^+e^-$
< 7	90	41	BESSON	03 CLE2	10.6 $e^+e^-$

### $D_{s1}(2460)^+$ DECAY MODES

$D_{s1}(2460)^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $D_s^{*+}\pi^0$	(48 ±11 ) %	S=1.1 CL=90%
$\Gamma_2$ $D_s^+\gamma$	(18 ± 4 ) %	
$\Gamma_3$ $D_s^{*+}\pi^-\pi^-$	( 4.3± 1.3 ) %	
$\Gamma_4$ $D_s^{*+}\gamma$	< 8 %	
$\Gamma_5$ $D_{s0}^+(2317)^+\gamma$	( 3.7 $^{+5.0}_{-2.4}$ ) %	
$\Gamma_6$ $D_s^+\pi^0$		
$\Gamma_7$ $D_s^+\pi^0\pi^0$		
$\Gamma_8$ $D_s^+\gamma\gamma$		

### CONSTRAINED FIT INFORMATION

An overall fit to 7 branching ratios uses 8 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 3.4$  for 4 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	80		
$x_3$	68	62	
$x_5$	−3	25	26
	$x_1$	$x_2$	$x_3$

### $D_{s1}(2460)^{\pm}$ BRANCHING RATIOS

$\Gamma(D_s^{*+}\pi^0)/\Gamma_{\text{total}}$	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>0.48±0.11 OUR FIT</b>					
<b>0.56±0.13±0.09</b>	<sup>11</sup>	AUBERT	06N BABR	$B \rightarrow D_{s1}(2460)^-\bar{D}^{(*)}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
	41	BESSON	03 CLE2	10.6 $e^+e^-$	
<sup>11</sup> Evaluated in AUBERT 06N including measurements from AUBERT,B 04s.					

$\Gamma(D_s^{*+}\gamma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
<b>0.18±0.04 OUR FIT</b>				
<b>0.16±0.04±0.03</b>	<sup>12</sup>	AUBERT	06N BABR	$B \rightarrow D_{s1}(2460)^-\bar{D}^{(*)}$
<sup>12</sup> Evaluated in AUBERT 06N including measurements from AUBERT,B 04s.				

$\Gamma(D_s^{*+}\gamma)/\Gamma(D_s^{*+}\pi^0)$	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_1$
<b>0.38 ±0.05 OUR FIT</b>						
<b>0.44 ±0.09 OUR AVERAGE</b>						
0.55 ±0.13 ±0.08	152	MIKAMI	04 BELL	10.6 $e^+e^-$		
0.38 ±0.11 ±0.04	38	KROKOVNY	03B BELL	10.6 $e^+e^-$		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.274 ±0.045 ±0.020	251	<sup>13</sup> AUBERT,B	04s BABR	$B \rightarrow D_{s1}(2460)^+\bar{D}^{(*)}$		
< 0.49	90	BESSON	03 CLE2	10.6 $e^+e^-$		
<sup>13</sup> Used by AUBERT 06N in their measurement of $B(D_s^{*-}\pi^0)$ and $B(D_s^-\gamma)$ .						

$\Gamma(D_s^{*+}\pi^+\pi^-)/\Gamma(D_s^{*+}\pi^0)$	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma_1$
<b>0.090±0.020 OUR FIT</b>	Error includes scale factor of 1.2.					
<b>0.14 ±0.04 ±0.02</b>	60	MIKAMI	04 BELL	10.6 $e^+e^-$		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.08	90	BESSON	03 CLE2	10.6 $e^+e^-$		

$\Gamma(D_s^{*+}\gamma)/\Gamma(D_s^{*+}\pi^0)$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma_1$
<b>&lt;0.16</b>	90	BESSON	03 CLE2	10.6 $e^+e^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.31	90	MIKAMI	04 BELL	10.6 $e^+e^-$	

$\Gamma(D_{s0}^+(2317)^+\gamma)/\Gamma(D_s^{*+}\pi^0)$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma_1$
<b>&lt;0.22</b>	95	AUBERT	04E BABR	10.6 $e^+e^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.58	90	BESSON	03 CLE2	10.6 $e^+e^-$	

$\Gamma(D_s^{*+}\pi^0)/[\Gamma(D_s^{*+}\pi^0) + \Gamma(D_{s0}^+(2317)^+\gamma)]$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/(\Gamma_1+\Gamma_5)$
<b>0.93±0.09 OUR FIT</b>				
<b>0.97±0.09±0.05</b>	AUBERT	06P BABR	10.6 $e^+e^-$	

$\Gamma(D_s^{*+}\gamma)/[\Gamma(D_s^{*+}\pi^0) + \Gamma(D_{s0}^+(2317)^+\gamma)]$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/(\Gamma_1+\Gamma_5)$
<b>0.35 ±0.04 OUR FIT</b>				
<b>0.337±0.036±0.038</b>	AUBERT	06P BABR	10.6 $e^+e^-$	

$\Gamma(D_s^{*+}\pi^+\pi^-)/[\Gamma(D_s^{*+}\pi^0) + \Gamma(D_{s0}^+(2317)^+\gamma)]$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/(\Gamma_1+\Gamma_5)$
<b>0.083±0.017 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.077±0.013±0.008</b>	AUBERT	06P BABR	10.6 $e^+e^-$	

$\Gamma(D_s^{*+}\gamma)/[\Gamma(D_s^{*+}\pi^0) + \Gamma(D_{s0}^+(2317)^+\gamma)]$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/(\Gamma_1+\Gamma_5)$
<b>&lt;0.24</b>	95	AUBERT	06P BABR	10.6 $e^+e^-$	

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## Meson Particle Listings

$$D_{s1}(2460)^{\pm}, D_{s1}(2536)^{\pm}$$

$\Gamma(D_{s0}^{*0}(2317)^+ \gamma) / [\Gamma(D_s^{*+} \pi^0) + \Gamma(D_{s0}^{*0}(2317)^+ \gamma)]$					$\Gamma_5 / (\Gamma_1 + \Gamma_5)$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.25	95	AUBERT	06P	BABR	10.6 $e^+ e^-$

$\Gamma(D_s^{*0} \pi^0) / [\Gamma(D_s^{*+} \pi^0) + \Gamma(D_{s0}^{*0}(2317)^+ \gamma)]$					$\Gamma_6 / (\Gamma_1 + \Gamma_5)$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.042	95	AUBERT	06P	BABR	10.6 $e^+ e^-$

$\Gamma(D_s^{*0} \pi^0 \pi^0) / [\Gamma(D_s^{*+} \pi^0) + \Gamma(D_{s0}^{*0}(2317)^+ \gamma)]$					$\Gamma_7 / (\Gamma_1 + \Gamma_5)$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.68	95	AUBERT	06P	BABR	10.6 $e^+ e^-$

$\Gamma(D_s^{*+} \gamma \gamma) / [\Gamma(D_s^{*+} \pi^0) + \Gamma(D_{s0}^{*0}(2317)^+ \gamma)]$					$\Gamma_8 / (\Gamma_1 + \Gamma_5)$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.33	95	AUBERT	06P	BABR	10.6 $e^+ e^-$

 $D_{s1}(2460)^{\pm}$  REFERENCES

AUBERT	06N	PR D74 031103	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	06P	PR D74 032007	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	04E	PR D69 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	04S	PRL 93 181801	B. Aubert <i>et al.</i>	(BABAR Collab.)
MIKAMI	04	PRL 92 012002	Y. Mikami <i>et al.</i>	(BELLE Collab.)
BESSION	03	PR D68 032002	D. Besson <i>et al.</i>	(CLEO Collab.)
KROKOVNY	03B	PRL 91 262002	P. Krokovny <i>et al.</i>	(BELLE Collab.)

$$D_{s1}(2536)^{\pm}$$

$$I(J^P) = 0(1^+)$$

J, P need confirmation.

Seen in  $D^*(2010)^+ K^0$ ,  $D^*(2007)^0 K^+$ , and  $D_s^{*+} \pi^+ \pi^-$ . Not seen in  $D^+ K^0$  or  $D^0 K^+$ .  $J^P = 1^+$  assignment strongly favored.

 $D_{s1}(2536)^{\pm}$  MASS

The fit includes  $D^{\pm}$ ,  $D^0$ ,  $D_s^{\pm}$ ,  $D^{*+}$ ,  $D^{*0}$ ,  $D_s^{*+}$ ,  $D_1(2420)^0$ ,  $D_2^*(2460)^0$ , and  $D_{s1}(2536)^{\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2535.10 ± 0.06 OUR FIT</b>				
<b>2535.18 ± 0.24 OUR AVERAGE</b>				
2535.7 ± 0.6 ± 0.5	46 ± 9	<sup>1</sup> ABAZOV	09G D0	$B_s^0 \rightarrow D_{s1}^- \mu^+ \nu_{\mu} X$
2534.78 ± 0.31 ± 0.40	182	AUBERT	08B BABR	$B \rightarrow \overline{D}^{(*)} D^{*0} K$
2534.6 ± 0.3 ± 0.7	193	AUBERT	06P BABR	10.6 $e^+ e^- \rightarrow D_s^{*+} \pi^+ \pi^- X$
2535.3 ± 0.7	92	<sup>2</sup> HEISTER	02B ALEP	$e^+ e^- \rightarrow D^{*+} K^0 X$ , $D^{*0} K^+ X$
2534.2 ± 1.2	9	ASRATYAN	94 BEBC	$\nu N \rightarrow D^{*0} K^0 X, D^{*0} K^{\pm} X$
2535 ± 0.6 ± 1	75	FRABETTI	94B E687	$\gamma \text{Be} \rightarrow D^{*+} K^0 X$ , $D^{*0} K^+ X$
2535.3 ± 0.2 ± 0.5	134	ALEXANDER	93 CLE2	$e^+ e^- \rightarrow D^{*0} K^+ X$
2534.8 ± 0.6 ± 0.6	44	ALEXANDER	93 CLE2	$e^+ e^- \rightarrow D^{*+} K^0 X$
2535.2 ± 0.5 ± 1.5	28	ALBRECHT	92R ARG	10.4 $e^+ e^- \rightarrow D^{*0} K^+ X$
2536.6 ± 0.7 ± 0.4		VERY	90 CLEO	$e^+ e^- \rightarrow D^{*+} K^0 X$
2535.9 ± 0.6 ± 2.0		ALBRECHT	89E ARG	$D_{s1}^+ \rightarrow D^*(2010) K^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2534.1 ± 0.6	116	<sup>3</sup> AUSHEV	11 BELL	$B \rightarrow D_{s1}(2536) + D^{(*)}$
2535.08 ± 0.01 ± 0.15	8038	<sup>4</sup> LEES	11B BABR	10.6 $e^+ e^- \rightarrow D^{*+} K_S^0 X$
2535.57 ± 0.44 ± 0.10	236 ± 30	<sup>5</sup> CHEKANOV	09 ZEUS	$e^{\pm} p \rightarrow D^{*+} K_S^0 X$ , $D^{*0} K^+ X$
2535 ± 28		<sup>6</sup> ASRATYAN	88 HLBC	$\nu N \rightarrow D_s \gamma \gamma X$

<sup>1</sup> Using the  $D^*(2010)^{\pm}$  mass of 2010.0 ± 0.4 MeV from PDG 06.

<sup>2</sup> Calculated using  $m(D^*(2010)^{\pm}) = 2010.0 \pm 0.5$  MeV,  $m(D^*(2007)^0) = 2006.7 \pm 0.5$  MeV, and the mass difference below.

<sup>3</sup> Systematic uncertainties not evaluated.

<sup>4</sup> Calculated using the mass difference  $m(D_{s1}^+) - m(D^{*+})_{PDG}$  below and  $m(D^{*+})_{PDG} = 2010.25 \pm 0.14$  MeV. Assuming S-wave decay of the  $D_{s1}(2536)$  to  $D^{*+} K_S^0$ , using a Breit-Wigner line shape corresponding to L=0.

<sup>5</sup> Calculated using the mass difference  $m(D_{s1}^+) - m(D^{*+})_{PDG}$  reported below and  $m(D^{*+})_{PDG} = 2010.27 \pm 0.17$  MeV.

<sup>6</sup> Not seen in  $D^* K$ .

$$m_{D_{s1}(2536)^{\pm}} - m_{D_s^{*0}(2111)}$$

The fit includes  $D^{\pm}$ ,  $D^0$ ,  $D_s^{\pm}$ ,  $D^{*+}$ ,  $D^{*0}$ ,  $D_s^{*+}$ ,  $D_1(2420)^0$ ,  $D_2^*(2460)^0$ , and  $D_{s1}(2536)^{\pm}$  mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>423.0 ± 0.4 OUR FIT</b>			
<b>424 ± 28</b>	ASRATYAN	88 HLBC	$D_s^{*+} \gamma$

$$m_{D_{s1}(2536)^{\pm}} - m_{D^{*0}(2010)^{\pm}}$$

The fit includes  $D^{\pm}$ ,  $D^0$ ,  $D_s^{\pm}$ ,  $D^{*+}$ ,  $D^{*0}$ ,  $D_s^{*+}$ ,  $D_1(2420)^0$ ,  $D_2^*(2460)^0$ , and  $D_{s1}(2536)^{\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>524.84 ± 0.04 OUR FIT</b>				
<b>524.84 ± 0.04 OUR AVERAGE</b>				
524.83 ± 0.01 ± 0.04	8038	<sup>7</sup> LEES	11B BABR	10.6 $e^+ e^- \rightarrow D^{*+} K_S^0 X$
525.30 ± 0.44 ± 0.10	236 ± 30	CHEKANOV	09 ZEUS	$e^{\pm} p \rightarrow D^{*+} K_S^0 X$ , $D^{*0} K^+ X$
525.3 ± 0.6 ± 0.1	41	HEISTER	02B ALEP	$e^+ e^- \rightarrow D^{*+} K^0 X$
<sup>7</sup> Assuming S-wave decay of the $D_{s1}(2536)$ to $D^{*+} K_S^0$ , using a Breit-Wigner line shape corresponding to L=0.				

$$m_{D_{s1}(2536)^{\pm}} - m_{D^{*0}(2007)^0}$$

The fit includes  $D^{\pm}$ ,  $D^0$ ,  $D_s^{\pm}$ ,  $D^{*+}$ ,  $D^{*0}$ ,  $D_s^{*+}$ ,  $D_1(2420)^0$ ,  $D_2^*(2460)^0$ , and  $D_{s1}(2536)^{\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>528.25 ± 0.05 OUR FIT</b>				Error includes scale factor of 1.1.
<b>528.1 ± 1.5 OUR AVERAGE</b>				
528.7 ± 1.9 ± 0.5	51	HEISTER	02B ALEP	$e^+ e^- \rightarrow D^{*0} K^+ X$
527.3 ± 2.2	29	ACKERSTAFF	97W OPAL	$e^+ e^- \rightarrow D^{*0} K^+ X$

 $D_{s1}(2536)^{\pm}$  WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.92 ± 0.03 ± 0.04</b>		8038	<sup>8</sup> LEES	11B BABR	10.6 $e^+ e^- \rightarrow D^{*+} K_S^0 X$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.75 ± 0.23		116	<sup>9</sup> AUSHEV	11 BELL	$B \rightarrow D_{s1}(2536) + D^{(*)}$
< 2.5	95	193	AUBERT	06P BABR	10.6 $e^+ e^- \rightarrow D_s^{*+} \pi^+ \pi^- X$
< 3.2	90	75	FRABETTI	94B E687	$\gamma \text{Be} \rightarrow D^{*+} K^0 X$ , $D^{*0} K^+ X$
< 2.3	90		ALEXANDER	93 CLEO	$e^+ e^- \rightarrow D^{*0} K^+ X$
< 3.9	90		ALBRECHT	92R ARG	10.4 $e^+ e^- \rightarrow D^{*0} K^+ X$
< 5.44	90		VERY	90 CLEO	$e^+ e^- \rightarrow D^{*+} K^0 X$
< 4.6	90		ALBRECHT	89E ARG	$D_{s1}^+ \rightarrow D^*(2010) K^0$

<sup>8</sup> Assuming S-wave decay of the  $D_{s1}(2536)$  to  $D^{*+} K_S^0$ , using a Breit-Wigner line shape corresponding to L=0.

<sup>9</sup> Systematic uncertainties not evaluated.

 $D_{s1}(2536)^+$  DECAY MODES

$D_{s1}(2536)^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i / \Gamma$ )	Confidence level
$\Gamma_1$ $D^*(2010)^+ K^0$	0.85 ± 0.12	
$\Gamma_2$ $(D^*(2010)^+ K^0)_{S\text{-wave}}$	0.61 ± 0.09	
$\Gamma_3$ $(D^*(2010)^+ K^0)_{D\text{-wave}}$		
$\Gamma_4$ $D^+ \pi^- K^+$	0.028 ± 0.005	
$\Gamma_5$ $D^*(2007)^0 K^+$	<b>DEFINED AS 1</b>	
$\Gamma_6$ $D^+ K^0$	<0.34	90%
$\Gamma_7$ $D^0 K^+$	<0.12	90%
$\Gamma_8$ $D_s^{*+} \gamma$	possibly seen	
$\Gamma_9$ $D_s^{*+} \pi^+ \pi^-$	seen	

 $D_{s1}(2536)^+$  BRANCHING RATIOS

$\Gamma(D^*(2007)^0 K^+) / \Gamma(D^*(2010)^+ K^0)$					$\Gamma_5 / \Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.18 ± 0.16 OUR AVERAGE</b>					
0.88 ± 0.24 ± 0.08	116	AUSHEV	11 BELL	$B \rightarrow D_{s1}(2536) + D^{(*)}$	
2.3 ± 0.6 ± 0.3	236 ± 30	CHEKANOV	09 ZEUS	$e^{\pm} p \rightarrow D^{*+} K_S^0 X$ , $D^{*0} K^+ X$	
1.32 ± 0.47 ± 0.23	92	<sup>10</sup> HEISTER	02B ALEP	$e^+ e^- \rightarrow D^{*+} K^0 X$ , $D^{*0} K^+ X$	
1.9 ± 1.1 ± 0.9	± 0.4	<sup>10</sup> ACKERSTAFF	97W OPAL	$e^+ e^- \rightarrow D^{*0} K^+ X$ , $D^{*+} K^0 X$	
1.1 ± 0.3		ALEXANDER	93 CLEO	$e^+ e^- \rightarrow D^{*0} K^+ X, D^{*+} K^0 X$	
1.4 ± 0.3 ± 0.2		<sup>11</sup> ALBRECHT	92R ARG	10.4 $e^+ e^- \rightarrow D^{*0} K^+ X, D^{*+} K^0 X$	

<sup>10</sup> Ratio of the production rates measured in  $Z^0$  decays.

<sup>11</sup> Evaluated by us from published inclusive cross-sections.

$\Gamma((D^*(2010)^+ K^0)_{S\text{-wave}}) / \Gamma(D^*(2010)^+ K^0)$					$\Gamma_2 / \Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.72 ± 0.05 ± 0.01</b>	5485	BALAGURA	08 BELL	10.6 $e^+ e^- \rightarrow D^{*+} K^0 X$	

# Meson Particle Listings

## $D_{s1}(2536)^\pm, D_{s2}(2573)$

$\Gamma(D^+\pi^-K^+)/\Gamma(D^*(2010)^+K^0)$					$\Gamma_4/\Gamma_1$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.27\pm 0.18\pm 0.37</math></b>	1264	BALAGURA	08	BELL	$10.6\,e^+e^-\rightarrow D^+\pi^-K^+X$
$\Gamma(D^+K^0)/\Gamma(D^*(2010)^+K^0)$					$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.40</b>	90	ALEXANDER	93	CLEO	$e^+e^-\rightarrow D^{*+}K^0X$
<0.43	90	ALBRECHT	89E	ARG	$D_{s1}^*\rightarrow D^*(2010)K^0$
$\Gamma(D^0K^+)/\Gamma(D^*(2007)^0K^+)$					$\Gamma_7/\Gamma_5$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.12</b>	90	ALEXANDER	93	CLEO	$e^+e^-\rightarrow D^{*0}K^+X$
$\Gamma(D_s^{*+}\gamma)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
possibly seen		ASRATYAN	88	HLBC	$\nu N\rightarrow D_s\gamma\gamma X$
$\Gamma(D_s^{*+}\gamma)/\Gamma(D^*(2007)^0K^+)$					$\Gamma_8/\Gamma_5$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.42</b>	90	ALEXANDER	93	CLEO	$e^+e^-\rightarrow D^{*0}K^+X$
$\Gamma(D_s^{*+}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
seen		AUBERT	06P	BABR	$10.6\,e^+e^-\rightarrow D_s^{*+}\pi^+\pi^-X$

### $D_{s1}(2536)^\pm$ REFERENCES

AUSHEV	11	PR D83 051102	T. Aushev <i>et al.</i>	(BELLE Collab.)
LEES	11B	PR D83 072003	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABAZOV	09G	PRL 102 051801	V.M. Abazov <i>et al.</i>	(D0 Collab.)
CHEKANOV	09	EPJ C60 25	S. Chekanov <i>et al.</i>	(ZEUS Collab.)
AUBERT	08B	PR D77 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
BALAGURA	08	PR D77 032001	V. Balagura <i>et al.</i>	(BELLE Collab.)
AUBERT	06P	PR D74 032007	B. Aubert <i>et al.</i>	(BABAR Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
HEISTER	02B	PL B526 34	A. Heister <i>et al.</i>	(ALEPH Collab.)
ACKERSTAFF	97W	ZPHY C76 425	K. Ackersstaff <i>et al.</i>	(OPAL Collab.)
ASRATYAN	94	ZPHY C61 563	A.E. Asratyan <i>et al.</i>	(BIRM, BELG, CERN+)
FRABETTI	94B	PRL 72 324	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALEXANDER	93	PL B303 377	J. Alexander <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92R	PL B297 425	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AVERY	90	PR D41 774	P. Avery, D. Besson	(CLEO Collab.)
ALBRECHT	89E	PL B230 162	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ASRATYAN	88	ZPHY C40 483	A.E. Asratyan <i>et al.</i>	(ITEP, SERP)

$D_{s2}^*(2573)$

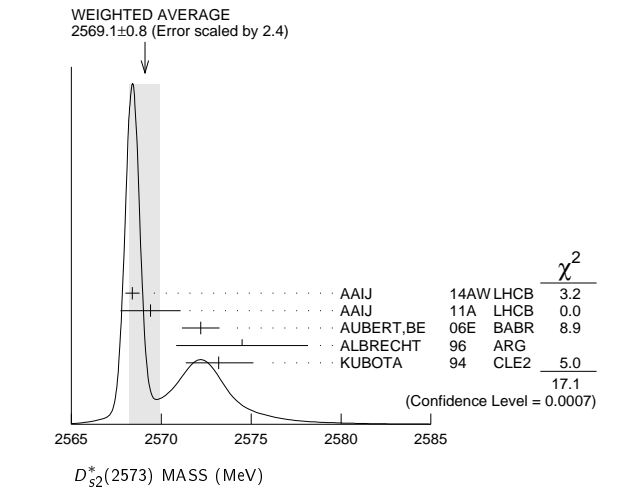
$I(J^P) = 0(2^+)$

$J^P$  is natural, width and decay modes consistent with  $2^+$ .  
AAIJ 14BJ confirms  $J^P = 2^+$ .

### $D_{s2}^*(2573)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2569.1\pm 0.8</math></b>	<b>OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.		
$2568.39\pm 0.29\pm 0.26$		AAIJ	14AW LHCB	$B_s^0\rightarrow\bar{D}^0K^-\pi^+$
$2569.4\pm 1.6\pm 0.5$	82	AAIJ	11A LHCB	$B_s\rightarrow D_{s2}^*(2573)\mu\bar{\nu}X$
$2572.2\pm 0.3\pm 1.0$		AUBERT,BE	06E BABR	$e^+e^-\rightarrow DKX$
$2574.5\pm 3.3\pm 1.6$		ALBRECHT	96 ARG	$e^+e^-\rightarrow D^0K^+X$
$2573.2\pm 1.7\pm 1.6$	217	KUBOTA	94 CLE2	$e^+e^-\sim 10.5\text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2570.0\pm 4.3$	25	<sup>1</sup> EVDOKIMOV	04 SELX	$600\,\Sigma^-A\rightarrow D^0K^+X$
$2568.6\pm 3.2$	64	<sup>2</sup> HEISTER	02B ALEP	$e^+e^-\rightarrow D^0K^+X$

<sup>1</sup> Not independent of the mass difference below.  
<sup>2</sup> Calculated using  $m_{D^0} = 1864.5 \pm 0.5$  MeV and the mass difference below.



$m_{D_{s2}^*(2573)} - m_{D^0}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>704\pm 3\pm 1</math></b>	64	HEISTER	02B ALEP	$e^+e^-\rightarrow D^0K^+X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$705.4\pm 4.3$	25	<sup>1</sup> EVDOKIMOV	04 SELX	$600\,\Sigma^-A\rightarrow D^0K^+X$
<sup>1</sup> Systematic errors not estimated.				

### $D_{s2}^*(2573)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>16.9\pm 0.8</math></b>	<b>OUR AVERAGE</b>			
$16.9\pm 0.5\pm 0.6$		AAIJ	14AW LHCB	$B_s^0\rightarrow\bar{D}^0K^-\pi^+$
$12.1\pm 4.5\pm 1.6$	82	AAIJ	11A LHCB	$B_s\rightarrow D_{s2}^*(2573)\mu\bar{\nu}X$
$27.1\pm 0.6\pm 5.6$		AUBERT,BE	06E BABR	$e^+e^-\rightarrow DKX$
$10.4\pm 8.3\pm 3.0$		ALBRECHT	96 ARG	$e^+e^-\rightarrow D^0K^+X$
$16\pm 5\pm 3$	217	KUBOTA	94 CLE2	$e^+e^-\sim 10.5\text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$14\pm 9\pm 6$	25	<sup>1</sup> EVDOKIMOV	04 SELX	$600\,\Sigma^-A\rightarrow D^0K^+X$
<sup>1</sup> Systematic errors not estimated.				

### $D_{s2}^*(2573)^+$ DECAY MODES

$D_{s2}^*(2573)^-$ modes are charge conjugates of the modes below.	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1\ D^0K^+$	seen
$\Gamma_2\ D^*(2007)^0K^+$	not seen

### $D_{s2}^*(2573)^+$ BRANCHING RATIOS

$\Gamma(D^0K^+)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
seen	217	KUBOTA	94 CLE2	$\pm$	$e^+e^-\sim 10.5\text{ GeV}$
$\Gamma(D^*(2007)^0K^+)/\Gamma(D^0K^+)$					$\Gamma_2/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;0.33</b>	90	KUBOTA	94 CLE2	+	$e^+e^-\sim 10.5\text{ GeV}$

### $D_{s2}^*(2573)$ REFERENCES

AAIJ	14AW	PRL 113 162001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14BJ	PRL 113 242002	R. Aaij <i>et al.</i>	(LHCb Collab.) JP
AAIJ	11A	PL B698 14	R. Aaij <i>et al.</i>	(LHCb Collab.)
AUBERT,BE	06E	PRL 97 222001	B. Aubert <i>et al.</i>	(BABAR Collab.)
EVDOKIMOV	04	PRL 93 242001	A.V. Evdokimov <i>et al.</i>	(SELEX Collab.)
HEISTER	02B	PL B526 34	A. Heister <i>et al.</i>	(ALEPH Collab.)
ALBRECHT	96	ZPHY C69 405	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
KUBOTA	94	PRL 72 1972	Y. Kubota <i>et al.</i>	(CLEO Collab.)

See key on page 601

## Meson Particle Listings

$$D_{s1}^*(2700)^\pm, D_{s1}^*(2860)^\pm$$

$$D_{s1}^*(2700)^\pm$$

$$I(J^P) = 0(1^-)$$

 $D_{s1}^*(2700)^+$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>2708.3<math>\pm</math> 4.0<math>\pm</math> 3.4</b>	<b>OUR AVERAGE</b>			
2699 $\pm$ 14 $\pm$ 7	52k	<sup>1</sup> LEES	15c BABR	$B \rightarrow D D^0 K^+$
2709.2 $\pm$ 1.9 $\pm$ 4.5	10.4k	<sup>2</sup> AAIJ	12AU LHCB	$pp \rightarrow (D K)^+ X$ at 7 TeV
2710 $\pm$ 2 $\pm$ 12 $\pm$ 7	182	<sup>3</sup> AUBERT	09AR BABR	$e^+ e^- \rightarrow D^{(*)} K X$
2708 $\pm$ 9 $\pm$ 11 $\pm$ 10		BRODZICKA	08 BELL	$B^+ \rightarrow D^0 \bar{D}^0 K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2694 $\pm$ 8 $\pm$ 13 $\pm$ 3		LEES	15c BABR	$B^0 \rightarrow D^- D^0 K^+$
2707 $\pm$ 8 $\pm$ 8		LEES	15c BABR	$B^+ \rightarrow \bar{D}^0 D^0 K^+$
2688 $\pm$ 4 $\pm$ 3		<sup>4</sup> AUBERT, BE	06E BABR	10.6 $e^+ e^- \rightarrow D K X$
<sup>1</sup> From a combined analysis of $B^0 \rightarrow D^- D^0 K^+$ and $B^+ \rightarrow \bar{D}^0 D^0 K^+$ .				
<sup>2</sup> From the combined fit of the $D^+ K_S^0$ and $D^0 K^+$ modes in the model including the $D_{s2}^*(2573)^+$ , $D_{s1}^*(2700)^+$ and spin-0 $D_{sJ}^*(2860)^+$ .				
<sup>3</sup> From simultaneous fits to the two $D K$ mass spectra and to the total $D^* K$ mass spectrum.				
<sup>4</sup> Superseded by AUBERT 09AR.				

 $D_{s1}^*(2700)^+$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>120<math>\pm</math> 11<math>\pm</math> 19</b>	<b>OUR AVERAGE</b>			
127 $\pm$ 24 $\pm$ 19	52k	<sup>5</sup> LEES	15c BABR	$B \rightarrow D D^0 K^+$
115.8 $\pm$ 7.3 $\pm$ 12.1	10.4k	<sup>6</sup> AAIJ	12AU LHCB	$pp \rightarrow (D K)^+ X$ at 7 TeV
149 $\pm$ 7 $\pm$ 39 $\pm$ 52	182	<sup>7</sup> AUBERT	09AR BABR	$e^+ e^- \rightarrow D^{(*)} K X$
108 $\pm$ 23 $\pm$ 36 $\pm$ 31		BRODZICKA	08 BELL	$B^+ \rightarrow D^0 \bar{D}^0 K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
145 $\pm$ 24 $\pm$ 22 $\pm$ 14		LEES	15c BABR	$B^0 \rightarrow D^- D^0 K^+$
113 $\pm$ 21 $\pm$ 20 $\pm$ 16		LEES	15c BABR	$B^+ \rightarrow \bar{D}^0 D^0 K^+$
112 $\pm$ 7 $\pm$ 36		<sup>8</sup> AUBERT, BE	06E BABR	10.6 $e^+ e^- \rightarrow D K X$
<sup>5</sup> From a combined analysis of $B^0 \rightarrow D^- D^0 K^+$ and $B^+ \rightarrow \bar{D}^0 D^0 K^+$ .				
<sup>6</sup> From the combined fit of the $D^+ K_S^0$ and $D^0 K^+$ modes in the model including the $D_{s2}^*(2573)^+$ , $D_{s1}^*(2700)^+$ and spin-0 $D_{sJ}^*(2860)^+$ .				
<sup>7</sup> From simultaneous fits to the two $D K$ mass spectra and to the total $D^* K$ mass spectrum.				
<sup>8</sup> Superseded by AUBERT 09AR.				

 $D_{s1}^*(2700)^\pm$  DECAY MODES

Mode
$\Gamma_1$ $D K$
$\Gamma_2$ $D^0 K^+$
$\Gamma_3$ $D^+ K_S^0$
$\Gamma_4$ $D^* K$
$\Gamma_5$ $D^{*0} K^+$
$\Gamma_6$ $D^{*+} K_S^0$

 $D_{s1}^*(2700)^\pm$  BRANCHING RATIOS

$\Gamma(D^* K)/\Gamma(D K)$	$\Gamma_4/\Gamma_1$
<b>0.91<math>\pm</math> 0.13<math>\pm</math> 0.12</b>	10.4k
<sup>9</sup> From the average of the corresponding ratios with $D^{(*)0} K^+$ and $D^{(*)+} K_S^0$ .	
$\Gamma(D^{*0} K^+)/\Gamma(D^0 K^+)$	$\Gamma_5/\Gamma_2$
<b>0.88<math>\pm</math> 0.14<math>\pm</math> 0.14</b>	7716
<sup>10</sup> From the $D^{*0} K^+$ and $D^0 K^+$ , where $D^{*0} \rightarrow D^0 \pi^0$ .	
$\Gamma(D^{*+} K_S^0)/\Gamma(D^+ K_S^0)$	$\Gamma_6/\Gamma_3$
<b>1.14<math>\pm</math> 0.39<math>\pm</math> 0.23</b>	2700
<sup>11</sup> From the $D^{*+} K_S^0$ and $D^+ K_S^0$ , where $D^{*+} \rightarrow D^+ \pi^0$ .	

 $D_{s1}^*(2700)^\pm$  REFERENCES

LEES	15c	PR D91 052002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
AAIJ	12AU	JHEP 1210 151	R. Aaij <i>et al.</i>	(LHCb Collab.)
AUBERT	09AR	PR D80 092003	B. Aubert <i>et al.</i>	(BABAR Collab.)
BRODZICKA	08	PRL 100 092001	J. Brodzicka <i>et al.</i>	(BELLE Collab.)
AUBERT, BE	06E	PRL 97 222001	B. Aubert <i>et al.</i>	(BABAR Collab.)

$$D_{s1}^*(2860)^\pm$$

$$I(J^P) = 0(1^-)$$

OMITTED FROM SUMMARY TABLE

$J^P$  consistent with  $1^-$  from angular analysis of AAIJ 14AW. Observed by AUBERT, BE 06E and AUBERT 09AR in inclusive production of  $D K$  and  $D^* K$  in  $e^+ e^-$  annihilation.

 $D_{s1}^*(2860)^+$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>2859<math>\pm</math> 12<math>\pm</math> 24</b>	<b>OUR AVERAGE</b>			
2866.1 $\pm$ 1.0 $\pm$ 6.3	36k	<sup>1,2,3</sup> AAIJ	14AW LHCB	$B_s^0 \rightarrow \bar{D}^0 K^- \pi^+$
2862 $\pm$ 2 $\pm$ 5 $\pm$ 2	3122	<sup>3,4</sup> AUBERT	09AR BABR	$pp \rightarrow (D K)^+ X$ at 7 TeV
2856.6 $\pm$ 1.5 $\pm$ 5.0		<sup>5</sup> AUBERT, BE	06E BABR	$e^+ e^- \rightarrow D K X$
<sup>1</sup> Separated from the spin-3 component $D_{s3}^*(2860)^-$ by a fit of the helicity angle of the $\bar{D}^0 K^-$ system, with a statistical significance of the spin-3 and spin-1 components in excess of $10 \sigma$ .				
<sup>2</sup> From the combined fit of the $D^+ K_S^0$ and $D^0 K^+$ modes in the model including the $D_{s2}^*(2573)^+$ , $D_{s1}^*(2700)^+$ and spin-0 $D_{sJ}^*(2860)^+$ .				
<sup>3</sup> Possible contribution from the $D_{s3}^*(2860)$ state.				
<sup>4</sup> From simultaneous fits to the two $D K$ mass spectra and to the total $D^* K$ mass spectrum.				
<sup>5</sup> Superseded by AUBERT 09AR.				

 $D_{s1}^*(2860)^+$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>159<math>\pm</math> 23<math>\pm</math> 77</b>	<b>OUR AVERAGE</b>			
69.9 $\pm$ 3.2 $\pm$ 6.6	36k	<sup>2,3</sup> AAIJ	12AU LHCB	$pp \rightarrow (D K)^+ X$ at 7 TeV
48 $\pm$ 3 $\pm$ 6	3122	<sup>3,4</sup> AUBERT	09AR BABR	$e^+ e^- \rightarrow D^{(*)} K X$
47 $\pm$ 7 $\pm$ 10		<sup>5</sup> AUBERT, BE	06E BABR	$e^+ e^- \rightarrow D K X$
<sup>1</sup> Separated from the spin-3 component $D_{s3}^*(2860)^-$ by a fit of the helicity angle of the $\bar{D}^0 K^-$ system, with a statistical significance of the spin-3 and spin-1 components in excess of $10 \sigma$ .				
<sup>2</sup> From the combined fit of the $D^+ K_S^0$ and $D^0 K^+$ modes in the model including the $D_{s2}^*(2573)^+$ , $D_{s1}^*(2700)^+$ and spin-0 $D_{sJ}^*(2860)^+$ .				
<sup>3</sup> Possible contribution from the $D_{s3}^*(2860)$ state.				
<sup>4</sup> From simultaneous fits to the two $D K$ mass spectra and to the total $D^* K$ mass spectrum.				
<sup>5</sup> Superseded by AUBERT 09AR.				

 $D_{s1}^*(2860)^\pm$  DECAY MODES

Mode
$\Gamma_1$ $D K$
$\Gamma_2$ $D^0 K^+$
$\Gamma_3$ $D^+ K_S^0$
$\Gamma_4$ $D^* K$
$\Gamma_5$ $D^{*0} K^+$
$\Gamma_6$ $D^{*+} K_S^0$

 $D_{s1}^*(2860)^\pm$  BRANCHING RATIOS

$\Gamma(D^* K)/\Gamma(D K)$	$\Gamma_4/\Gamma_1$
<b>1.10<math>\pm</math> 0.15<math>\pm</math> 0.19</b>	3122
<sup>1</sup> From the average of the corresponding ratios with $D^{(*)0} K^+$ and $D^{(*)+} K_S^0$ .	
$\Gamma(D^{*0} K^+)/\Gamma(D^0 K^+)$	$\Gamma_5/\Gamma_2$
<b>1.04<math>\pm</math> 0.17<math>\pm</math> 0.20</b>	2241
<sup>1</sup> From the $D^{*0} K^+$ and $D^0 K^+$ , where $D^{*0} \rightarrow D^0 \pi^0$ .	

Meson Particle Listings

$D_{s1}^*(2860)^\pm, D_{sJ}(3040)^\pm$

$\Gamma(D^{*+}K_S^0)/\Gamma(D^+K_S^0)$					$\Gamma_6/\Gamma_3$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.38 \pm 0.35 \pm 0.49$	881	<sup>1</sup> AUBERT	09AR BABR	$e^+e^- \rightarrow D^{(*)}KX$	
<sup>1</sup> From the $D^{*+}K_S^0$ and $D^+K_S^0$ , where $D^{*+} \rightarrow D^+\pi^0$ .					

$D_{s1}^*(2860)^\pm$ REFERENCES				
AAIJ	14AW PRL 113 162001	R. Aaij <i>et al.</i>	(LHCb Collab.)JP	
AAIJ	12AU JHEP 1210 151	R. Aaij <i>et al.</i>	(LHCb Collab.)	
AUBERT	09AR PR D80 092003	B. Aubert <i>et al.</i>	(BABAR Collab.)	
AUBERT,BE	06E PRL 97 222001	B. Aubert <i>et al.</i>	(BABAR Collab.)	

$D_{s3}^*(2860)^\pm$	$I(J^P) = 0(3^-)$
OMITTED FROM SUMMARY TABLE	
$J^P$ consistent with $3^-$ from angular analysis of AAIJ 14AW.	

$D_{s3}^*(2860)^+$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$2860.5 \pm 2.6 \pm 6.5$	<sup>1</sup> AAIJ	14AW LHCb	$B_S^0 \rightarrow \overline{D}^0 K^- \pi^+$
<sup>1</sup> Separated from the spin-1 component $D_{s1}^*(2860)^-$ by a fit of the helicity angle of the $\overline{D}^0 K^-$ system, with a statistical significance of the spin-3 and spin-1 components in excess of 10 $\sigma$ .			

$D_{s3}^*(2860)^+$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$53 \pm 7 \pm 7$	<sup>1</sup> AAIJ	14AW LHCb	$B_S^0 \rightarrow \overline{D}^0 K^- \pi^+$
<sup>1</sup> Separated from the spin-1 component $D_{s1}^*(2860)^-$ by a fit of the helicity angle of the $\overline{D}^0 K^-$ system, with a statistical significance of the spin-3 and spin-1 components in excess of 10 $\sigma$ .			

$D_{s3}^*(2860)^\pm$ REFERENCES			
AAIJ	14AW PRL 113 162001	R. Aaij <i>et al.</i>	(LHCb Collab.)JP

$D_{sJ}(3040)^\pm$	$I(J^P) = 0(?^?)$
OMITTED FROM SUMMARY TABLE	
Observed by AUBERT 09AR in inclusive production of $D^*K$ in $e^+e^-$ annihilation.	

$D_{sJ}(3040)^+$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$3044 \pm 8 \pm 30 \pm 5$	AUBERT	09AR BABR	$e^+e^- \rightarrow D^*KX$

$D_{sJ}(3040)^+$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$239 \pm 35 \pm 46 \pm 42$	AUBERT	09AR BABR	$e^+e^- \rightarrow D^*KX$

$D_{sJ}(3040)^\pm$ DECAY MODES	
Mode	
$\Gamma_1$	$D^*K$
$\Gamma_2$	$D^{*0}K^+$
$\Gamma_3$	$D^{*+}K_S^0$

$D_{sJ}(3040)^\pm$ REFERENCES			
AUBERT	09AR PR D80 092003	B. Aubert <i>et al.</i>	(BABAR Collab.)
OTHER RELATED PAPERS			
SUN	09 PR D80 074037	Z.-F. Sun, X. Lin	



**BOTTOM MESONS**

$$(B = \pm 1)$$

$$B^+ = u\bar{b}, B^0 = d\bar{b}, \bar{B}^0 = \bar{d}b, B^- = \bar{u}b, \text{ similarly for } B^{*'}s$$

***B*-particle organization**

Many measurements of *B* decays involve admixtures of *B* hadrons. Previously we arbitrarily included such admixtures in the  $B^\pm$  section, but because of their importance we have created two new sections: “ $B^\pm/B^0$  Admixture” for  $T(4S)$  results and “ $B^\pm/B^0/B_s^0/b$ -baryon Admixture” for results at higher energies. Most inclusive decay branching fractions and  $\chi_b$  at high energy are found in the Admixture sections.  $B^0\text{--}\bar{B}^0$  mixing data are found in the  $B^0$  section, while  $B_s^0\text{--}\bar{B}_s^0$  mixing data and  $B\text{--}\bar{B}$  mixing data for a  $B^0/B_s^0$  admixture are found in the  $B_s^0$  section. *CP*-violation data are found in the  $B^\pm$ ,  $B^0$ , and  $B^\pm/B^0$  Admixture sections. *b*-baryons are found near the end of the Baryon section. Recently, we also created a new section: “ $V_{cb}$  and  $V_{ub}$  CKM Matrix Elements.”

The organization of the *B* sections is now as follows, where bullets indicate particle sections and brackets indicate reviews.

[Production and Decay of *b*-flavored Hadrons]

[A Short Note on HFAG Activities]

- $B^\pm$ 
  - mass, mean life
  - branching fractions
  - polarization in  $B^\pm$  decay
  - CP* violation
- $B^0$ 
  - mass, mean life
  - branching fractions
  - [Polarization in *B* decay]
  - polarization in  $B^0$  decay
  - [*B*– $\bar{B}$  Mixing]
  - $B^0\text{--}\bar{B}^0$  mixing
  - CP* violation
- $B^\pm/B^0$  Admixture
  - branching fractions, *CP* violation
  - CP* violation
- $B^\pm/B^0/B_s^0/b$ -baryon Admixture
  - mean life
  - production fractions
  - branching fractions
  - $\chi_b$  at high energy
  - production fractions in hadronic *Z* decay
- $V_{cb}$  and  $V_{ub}$  CKM Matrix Elements
  - [Determination of  $V_{cb}$  and  $V_{ub}$ ]
- $B^*$ 
  - mass
- $B_1(5721)^0$ 
  - mass
- $B_J^*(5732)$ 
  - mass, width
- $B_2(5747)^0$ 
  - mass
- $B_s^0$ 
  - mass, mean life
  - branching fractions
  - polarization in  $B_s^0$  decay
  - $B_s^0\text{--}\bar{B}_s^0$  mixing
- $B_s^*$ 
  - mass
- $B_{s,J}^*(5850)$ 
  - mass, width
- $B_c^\pm$ 
  - mass, mean life
  - branching fractions

At the end of Baryon Listings:

- $\Lambda_b$ 
  - mass, mean life
  - branching fractions
- $\Sigma_b, \Sigma_b^*$ 
  - mass
- $\Xi_b^0, \Xi_b^-$ 
  - mean life
- $\Omega_b^-$ 
  - mass, mean life
  - branching fractions
- *b*-baryon Admixture
  - mean life
  - branching fractions

**PRODUCTION AND DECAY OF *b*-FLAVORED HADRONS**

Updated May 2016 by P. Eerola (U. of Helsinki, Helsinki, Finland), M. Kreps (U. of Warwick, Coventry, UK) and Y. Kwon (Yonsei U., Seoul, Korea).

The *b* quark belongs to the third generation of quarks and is the weak-doublet partner of the *t* quark. The existence of the third-generation quark doublet was proposed in 1973 by Kobayashi and Maskawa [1] in their model of the quark mixing matrix (“CKM” matrix), and confirmed four years later by the first observation of a  $b\bar{b}$  meson [2]. In the KM model, *CP* violation is explained within the Standard Model (SM) by an irreducible phase of the  $3 \times 3$  unitary matrix. The regular pattern of the three lepton and quark families is one of the most intriguing puzzles in particle physics. The existence of families gives rise to many of the free parameters in the SM, including the fermion masses, and the elements of the CKM matrix.

Since the *b* quark is the lighter element of the third-generation quark doublet, the decays of *b*-flavored hadrons occur via generation-changing processes through this matrix. Because of this, and the fact that the CKM matrix is close to a  $3 \times 3$  unit matrix, many interesting features such as loop and box diagrams, flavor oscillations, as well as large *CP* asymmetries, can be observed in the weak decays of *b*-flavored hadrons.

The CKM matrix is parameterized by three real parameters and one complex phase. This complex phase can become a source of *CP* violation in *B* meson decays. A crucial milestone was the first observation of *CP* violation in the *B* meson system in 2001, by the BaBar [3] and Belle [4] collaborations. They measured a large value for the parameter  $\sin 2\beta$  ( $= \sin 2\phi_1$ ) [5], almost four decades after the discovery of a small *CP* asymmetry in neutral kaons. A more detailed discussion of the CKM matrix and *CP* violation can be found elsewhere in this *Review* [6,7].

Recent developments in the physics of *b*-hadrons include the significant improvement in experimental determination of the CKM angle  $\gamma$ , the increased information on  $B_s$ ,  $B_c$  and  $\Lambda_b$  decays, the precise determination of  $\Lambda_b$  lifetime, the wealth of information in the  $B^0 \rightarrow K^{*0}(892)\ell^+\ell^-$  decays and after many years of search, the observation of  $B_s \rightarrow \mu^+\mu^-$  decays along with ever increasing precision on the CKM matrix parameters.

# Meson Particle Listings

## *b*-flavored hadrons

The structure of this mini-review is organized as follows. After a brief description of theory and terminology, we discuss *b*-quark production and current results on spectroscopy and lifetimes of *b*-flavored hadrons. We then discuss some basic properties of *B*-meson decays, followed by summaries of hadronic, rare, and electroweak penguin decays of *B*-mesons. There are separate mini-reviews for  $B\bar{B}$  mixing [8] and the extraction of the CKM matrix elements  $V_{cb}$  and  $V_{ub}$  from *B*-meson decays [9] in this *Review*.

**Theory and terminology:** The ground states of *b*-flavored hadrons decay via weak interactions. In most hadrons, the *b*-quark is accompanied by light-partner quarks (*d*, *u*, or *s*), and the decay modes are well described by the decay of the *b* quark (spectator model) [10]. The dominant decay mode of a *b* quark is  $b \rightarrow cW^{*-}$  (referred to as a “tree” or “spectator” decay), where the virtual *W* materializes either into a pair of leptons  $\ell\bar{\nu}$  (“semileptonic decay”), or into a pair of quarks which then hadronizes. The decays in which the spectator quark combines with one of the quarks from  $W^*$  to form one of the final state hadrons are suppressed by a factor  $\sim (1/3)^2$ , because the colors of the two quarks from different sources must match (“color-suppression”).

Many aspects of *B* decays can be understood through the Heavy Quark Effective Theory (HQET) [11]. This has been particularly successful for semileptonic decays. For further discussion of HQET, see for instance Ref. 12. For hadronic decays, one typically uses effective Hamiltonian calculations that rely on a perturbative expansion with Wilson coefficients. In addition, some form of the factorization hypothesis is commonly used, where, in analogy with semileptonic decays, two-body hadronic decays of *B* mesons are expressed as the product of two independent hadronic currents, one describing the formation of a charm meson (in case of the dominant  $b \rightarrow cW^{*-}$  decays), and the other the hadronization of the remaining  $\bar{u}d$  (or  $\bar{c}s$ ) system from the virtual  $W^-$ . Qualitatively, for a *B* decay with a large energy release, the  $\bar{u}d$  pair (produced as a color singlet) travels fast enough to leave the interaction region without influencing the charm meson. This is known to work well for the dominant spectator decays [13]. There are several common implementations of these ideas for hadronic *B* decays, the most common of which are QCD factorization (QCDF) [14], perturbative QCD (pQCD) [15], and soft collinear effective theory (SCET) [16].

The transition  $b \rightarrow u$  is suppressed by  $|V_{ub}/V_{cb}|^2 \sim (0.1)^2$  relative to  $b \rightarrow c$  transitions. The transition  $b \rightarrow s$  is a flavor-changing neutral-current (FCNC) process, and although not allowed in the SM as a tree-process, can occur via more complex loop diagrams (denoted “penguin” decays). The rates for such processes are comparable or larger than CKM-suppressed  $b \rightarrow u$  processes. Penguin processes involving  $b \rightarrow d$  transitions are also possible, and have been observed [17,18]. Other decay processes discussed in this *Review* include *W*-exchange (a *W* is exchanged between initial-state quarks), penguin annihilation (the gluon from a penguin loop attaches to the spectator quark,

similar to an exchange diagram), and pure-annihilation (the initial quarks annihilate to a virtual *W*, which then decays).

**Production and spectroscopy:** The bound states of a  $\bar{b}$  antiquark and a *u*, *d*, *s*, or *c* quark are referred to as the  $B_u$  ( $B^+$ ),  $B_d$  ( $B^0$ ),  $B_s$ , and  $B_c$  mesons, respectively. The  $B_c$  is the heaviest of the ground-state *b*-flavored mesons, and the most difficult to produce: it was observed for the first time in the semileptonic mode by CDF in 1998 [19], but its mass was accurately determined only in 2006, from the fully reconstructed mode  $B_c^+ \rightarrow J/\psi\pi^+$  [20]. One of the best determination up to date uses  $B_c^+ \rightarrow J/\psi D_s^+$  decay and yields  $m(B_c^+) = 6276.28 \pm 1.44 \pm 0.36$  MeV/ $c^2$  [21]. As this decay has very low energy release, it allows to decrease systematic uncertainty and thus offers prospects for future increase in precision.

The first excited meson is called the  $B^*$  meson, while  $B^{**}$  is the generic name for the four orbitally excited ( $L = 1$ ) *B*-meson states that correspond to the *P*-wave mesons in the charm system,  $D^{**}$ . Excited states of the  $B_s$  meson are similarly named  $B_s^*$  and  $B_s^{**}$ . Of the possible bound  $\bar{b}b$  states, the  $\Upsilon$  series (S-wave) and the  $\chi_b$  (P-wave) are well studied. The pseudoscalar ground state  $\eta_b$  also has been observed by BaBar [22] (and confirmed by CLEO [23]), indirectly through the decay  $\Upsilon(3S) \rightarrow \gamma\eta_b$ . See Ref. 24 for classification and naming of these and other states.

Experimental studies of *b* decays have been performed in  $e^+e^-$  collisions at the  $\Upsilon(4S)$  (ARGUS, CLEO, Belle, BaBar) and  $\Upsilon(5S)$  (CLEO, Belle) resonances, as well as at higher energies, at the *Z* resonance (SLC, LEP), in  $p\bar{p}$  (Tevatron) and  $pp$  collisions (LHC). The  $e^+e^- \rightarrow \bar{b}b$  production cross-section at the *Z*,  $\Upsilon(4S)$ , and  $\Upsilon(5S)$  resonances are about 6.6 nb, 1.1 nb, and 0.3 nb respectively. High-energy hadron collisions produce *b*-flavored hadrons of all species with much larger cross-sections:  $\sigma(p\bar{p} \rightarrow bX, |\eta| < 1) \sim 30$   $\mu\text{b}$  at the Tevatron ( $\sqrt{s} = 1.96$  TeV), and even higher at the energies of the LHC  $pp$  collider (at  $\sqrt{s} = 7$  TeV, at the LHCb experiment with pseudorapidity acceptance  $2 < \eta < 5$  visible *b*-hadron cross section is  $\sim 100$   $\mu\text{b}$ ).

BaBar and Belle have accumulated respectively 560 fb $^{-1}$  and 1020 fb $^{-1}$  of data, of which 433 fb $^{-1}$  and 710 fb $^{-1}$  respectively are at the  $\Upsilon(4S)$  resonance; CDF and D0 have accumulated by the end of their running about 10 fb $^{-1}$  each. At the LHC, CMS and ATLAS have collected 5 fb $^{-1}$  (20 fb $^{-1}$ ) of data at  $\sqrt{s} = 7$  (8) TeV respectively and LHCb has collected about 1 fb $^{-1}$  and 2 fb $^{-1}$  at the two energies. Further data was collected at  $\sqrt{13}$  TeV, but amount is limited at the moment. These numbers indicate that the majority of *b*-quarks have been produced in hadron collisions, but the large backgrounds cause the hadron collider experiments to have lower selection efficiency. While traditionally only the few decay modes for which triggering and reconstruction are easiest have been studied in hadron collisions, with current experiments at hadron colliders much more is possible. This is due to triggers based

on the tracking first introduced in CDF and further improved by LHCb. LHCb experiment has also reasonable capability for detection of neutral pions and photons. While both  $e^+e^-$  and hadron colliders have their own strengths and weaknesses, in the domain of decays which involve neutrinos,  $e^+e^-$  experiments are in significant advantage.

In hadron collisions, most production happens as  $b\bar{b}$  pairs, either via  $s$ -channel production or gluon-splitting, with a smaller fraction of single  $b$ -quarks produced by flavor excitation. The total  $b$ -production cross section is an interesting test of our understanding of QCD processes. For many years, experimental measurements have been several times higher than predictions. With improved measurements [25], more accurate input parameters, and more advanced calculations [26], the discrepancy between theory and data diminished and there is now good agreement between measurements and predictions.

Each quark of a  $b\bar{b}$  pair produced in hadron collisions hadronizes separately and incoherently from the other, but it is still possible, although difficult, to obtain a statistical indication of the charge of a produced  $b/\bar{b}$  quark (“flavor tag” or “charge tag”) from the accompanying particles produced in the hadronization process, or from the decay products of the other quark. The momentum spectrum of produced  $b$ -quarks typically peaks near the  $b$ -quark mass, and extends to much higher momenta, dropping by about a decade for every ten GeV. This implies typical decay lengths of the order of a millimeter; the resolution for the decay vertex must be more precise than this to resolve the fast oscillations of  $B_s$  mesons.

In  $e^+e^-$  colliders, since the  $B$  mesons are very slow in the  $\Upsilon(4S)$  rest frame, asymmetric beam energies are used to boost the decay products to improve the precision of time-dependent measurements that are crucial for the study of  $CP$  violation. At KEKB, the boost is  $\beta\gamma = 0.43$ , and the typical  $B$ -meson decay length is dilated from  $\approx 20 \mu\text{m}$  to  $\approx 200 \mu\text{m}$ . PEP-II used a slightly larger boost,  $\beta\gamma = 0.55$ . The two  $B$  mesons produced in  $\Upsilon(4S)$  decay are in a coherent quantum state, which makes it easier than in hadron collisions to infer the charge state of one  $B$  meson from observation of the other; however, the coherence also requires determination of the decay time of both mesons, rather than just one, in order to perform time-dependent  $CP$ -violation measurements. For  $B_s$ , which can be produced at  $\Upsilon(5S)$  the situation is less favourable, as boost is not high enough to provide sufficient time resolution to resolve the fast  $B_s$  oscillations.

For the measurement of branching fractions, the initial composition of the data sample must be known. The  $\Upsilon(4S)$  resonance decays predominantly to  $B^0\bar{B}^0$  and  $B^+B^-$ ; the current experimental upper limit for non- $B\bar{B}$  decays of the  $\Upsilon(4S)$  is less than 4% at the 95% confidence level (CL) [27]. The only known modes of this category are decays to lower  $\Upsilon$  states and a pion pair, observed with branching fractions of order  $10^{-4}$  [28]. The ratio  $f_+/f_0$  of the fractions of charged to neutral  $B$  productions from  $\Upsilon(4S)$  decays has been measured by CLEO, BaBar, and Belle in various ways. They typically

use pairs of isospin-related decays of  $B^+$  and  $B^0$ , such that it can be assumed that  $\Gamma(B^+ \rightarrow x^+) = \Gamma(B^0 \rightarrow x^0)$ . In this way, the ratio of the number of events observed in these modes is proportional to  $(f_+\tau_+)/ (f_0\tau_0)$  [29,30]. BaBar has also performed an independent measurement of  $f_0$  with a different method that does not require isospin symmetry or the value of the lifetime ratio, based on the number of events with one or two reconstructed  $B^0 \rightarrow D^{*-}\ell^+\nu$  decays [31]. The combined result, from the current average of  $\tau_+/\tau_0$ , is  $f_+/f_0 = 1.059 \pm 0.027$  [32]. Though the current  $2.4\sigma$  discrepancy with equal production of  $B^+B^-$  and  $B^0\bar{B}^0$  pairs is somewhat larger than previous averages, we still assume  $f_+/f_0 = 1$  in this mini-review except where explicitly stated otherwise. This assumption is also supported by the near equality of the  $B^+$  and  $B^0$  masses: our fit of CLEO, ARGUS, CDF, and LHCb measurements yields  $m(B^0) = 5279.61 \pm 0.16 \text{ MeV}/c^2$ ,  $m(B^+) = 5279.29 \pm 0.15 \text{ MeV}/c^2$ , and  $m(B^0) - m(B^+) = 0.32 \pm 0.06 \text{ MeV}/c^2$ .

CLEO and Belle have also collected some data at the  $\Upsilon(5S)$  resonance [33,34]. Belle has accumulated more than  $120 \text{ fb}^{-1}$  at this resonance. This resonance does not provide the simple final states like the  $\Upsilon(4S)$ : there are seven possible final states with a pair of non-strange  $B$  mesons and three with a pair of strange  $B$  mesons ( $B_s^*\bar{B}_s^*$ ,  $B_s^*\bar{B}_s$ , and  $B_s\bar{B}_s^*$ ). The fraction of events with a pair of  $B_s$  mesons over the total number of events with a pair of  $b$ -flavored hadrons has been measured to be  $f_s[\Upsilon(5S)] = 0.200^{+0.030}_{-0.031}$ , of which 90% is  $B_s^*\bar{B}_s^*$  events. A few branching fractions of the  $B_s$  have been measured in this way; if the precision of  $f_s$  were improved, they would become the most accurate. Belle has observed a few new  $B_s$  modes that are difficult to reconstruct in hadron colliders and the most precise mass measurement of the  $B_s^*$  meson has been obtained [34]. However, the small boost of  $B_s$  mesons produced in this way prevents resolution of their fast oscillations for time-dependent measurements; these are only accessible in hadron collisions or at the  $Z$  peak.

In high-energy collisions, the produced  $b$  or  $\bar{b}$  quarks can hadronize with different probabilities into the full spectrum of  $b$ -hadrons, either in their ground or excited states. Table 1 shows the measured fractions  $f_d$ ,  $f_u$ ,  $f_s$ , and  $f_{\text{baryon}}$  of  $B^0$ ,  $B^+$ ,  $B_s^0$ , and  $b$  baryons, respectively, in an unbiased sample of weakly decaying  $b$  hadrons produced at the  $Z$  resonance or in  $p\bar{p}$  collisions [32]. The results were obtained from a fit where the sum of the fractions were constrained to equal 1.0, neglecting production of  $B_c$  mesons. The observed yields of  $B_c$  mesons at the Tevatron [19] yields  $f_c = 0.2\%$ , in agreement with expectations [35], and well below the current experimental uncertainties in the other fractions.

For rather long time, the average of fractions in  $p\bar{p}$  collisions and in  $Z$  decay was used as it was assumed that the hadronization is identical in the two environments. It was clear that this assumption does not have to hold in principle, because of the different momentum distributions of the  $b$ -quark in these processes; the sample used in the  $p\bar{p}$  measurements has momenta close to the  $b$  mass, rather than  $m_Z/2$ . But

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## $b$ -flavored hadrons

**Table 1:** Fragmentation fractions of  $b$  quarks into weakly-decaying  $b$ -hadron species in  $Z \rightarrow b\bar{b}$  decay, in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV.

$b$ hadron	Fraction at $Z$ [%]	Fraction at $p\bar{p}$ [%]
$B^+, B^0$	$40.7 \pm 0.7$	$34.4 \pm 2.1$
$B_s$	$10.0 \pm 0.8$	$11.5 \pm 1.3$
$b$ baryons	$8.5 \pm 1.1$	$19.7 \pm 4.6$

in the absence of any significant evidence there was also no strong reason against the average. Some discrepancies were observed, but as picture was also obscured by  $1.8\sigma$  discrepancy in the average time-integrated mixing probability parameter  $\bar{\chi} = f_d\chi_d + f_s\chi_s$  between LEP and Tevatron [8], they were not directly attributed to breakdown of the assumption that hadronization is identical. The first indication that fraction for  $b$ -baryons depends on the momentum and thus environment came from CDF [36], but available precision did not allow for firm conclusion. The final evidence for non-universality of hadronization fractions came from LHCb, where strong dependence on the transverse momentum was observed for the  $\Lambda_b$  fraction [37].

Excited  $B$ -meson states have been observed by CLEO, LEP, CUSB, D0, and CDF. The current world average of the  $B^*-B$  mass difference is  $45.78 \pm 0.35$  MeV/ $c^2$ . Evidence for  $B^{**}$  ( $L=1$ ) production has been initially obtained at LEP [38], as a broad resonance in the mass of an inclusively reconstructed bottom hadron candidate combined with a charged pion from the primary vertex. Detailed results from exclusive modes have been obtained at the Tevatron, allowing separation of the narrow states  $B_1$  and  $B_2^*$  and also a measurement of the  $B_2^*$  width [39].

Also the narrow  $B_s^{**}$  states, first sighted by OPAL as a single broad enhancement in the  $B^+K$  mass spectrum [40], have now been clearly observed and separately measured at the hadron colliders [41,42]. The measured masses are  $m(B_{s1}) = 5828.7 \pm 0.4$  MeV/ $c^2$  and  $m(B_{s2}^*) = 5839.96 \pm 0.2$  MeV/ $c^2$ .

Baryon states containing a  $b$  quark are labeled according to the same scheme used for non- $b$  baryons, with the addition of a  $b$  subscript [24]. For many years, the only well-established  $b$  baryon was the  $\Lambda_b^0$  (quark composition  $udb$ ), with only indirect evidence for  $\Xi_b$  ( $dsb$ ) production from LEP [43]. This situation has changed dramatically in the past few years due to the large samples being accumulated at the Tevatron and LHCb. Clear signals of four strongly-decaying baryon states,  $\Sigma_b^+, \Sigma_b^{*+}$  ( $uub$ ),  $\Sigma_b^-, \Sigma_b^{*-}$  ( $ddb$ ) have been obtained by CDF in  $\Lambda_b^0\pi^\pm$  final states [44]. The strange bottom baryon  $\Xi_b^\pm$  was observed in the exclusive mode  $\Xi_b^\pm \rightarrow J/\psi\Xi^\pm$  by D0 [45], and CDF [46]. More recently CDF has also observed the  $\Xi_b$  in the  $\Xi_c\pi$  final state [47]. The relative production of  $\Xi_b$  and  $\Lambda_b$  baryons has been found to be consistent with the  $B_s$  to  $B_d$  production ratio [45]. Observation of the doubly-strange bottom baryon  $\Omega_b^-$  has been published by both D0 [48] and CDF [49]. However the masses

measured by the two experiments show a large discrepancy. The resolution is provided by LHCb which measures the  $\Omega_b^-$  mass consistent with CDF [50]. The CMS experiment added to the list also neutral spin-3/2  $\Xi_b^*$  [51]. The masses of all these new baryons have been measured to a precision of a few MeV/ $c^2$ , and found to be in agreement with predictions from HQET.

While many exotic states were seen in the charm sector, in bottom sector there are fewer seen and none in the direct production. In the recent analysis D0 Collaboration claimed narrow state  $X(5568)$  decaying into  $B_s^0\pi^\pm$  final state [52]. While this would be interesting addition to the observed states as first exotic state with open heavy flavour quantum numbers, preliminary analysis from LHCb yields negative result [53]. While two experiments have different initial state, it would be unusual if  $X(5568)$  can be effectively produced in high energy  $p\bar{p}$  collisions but not in  $pp$  collisions.

**Lifetimes:** Precise lifetimes are key in extracting the weak parameters that are important for understanding the role of the CKM matrix in  $CP$  violation, such as the determination of  $V_{cb}$  and  $B_s\bar{B}_s$  mixing parameters. In the naive spectator model, the heavy quark can decay only via the external spectator mechanism, and thus, the lifetimes of all mesons and baryons containing  $b$  quarks would be equal. Non-spectator effects, such as the interference between contributing amplitudes, modify this simple picture and give rise to a lifetime hierarchy for  $b$ -flavored hadrons similar to the one in the charm sector. However, since the lifetime differences are expected to scale as  $1/m_Q^2$ , where  $m_Q$  is the mass of the heavy quark, the variations in the  $b$  system are expected to be only 10% or less [54]. We expect:

$$\tau(B^+) \geq \tau(B^0) \approx \tau(B_s) > \tau(\Lambda_b^0) \gg \tau(B_c^+) . \quad (1)$$

For the  $B_c^+$ , both quarks decay weakly, so the lifetime is much shorter.

Measurements of the lifetimes of the different  $b$ -flavored hadrons thus provide a means to determine the importance of non-spectator mechanisms in the  $b$  sector. Over the past decade, the precision of silicon vertex detectors and the increasing availability of fully-reconstructed samples has resulted in much-reduced statistical and systematic uncertainties ( $\sim 1\%$ ). The averaging of precision results from different experiments is a complex task that requires careful treatment of correlated systematic uncertainties; the world averages given in Table 2 have been determined by the Heavy Flavor Averaging Group (HFAG) [32].

The short  $B_c^+$  lifetime is in good agreement with predictions [55]. With large samples of  $B_c^+$  mesons at the LHCb precision on the lifetimes should significantly improve. The measurement using semileptonic decays gives  $\tau_{B_c^+} = 0.509 \pm 0.008 \pm 0.012$  ps [56] while using decays  $B_c^+ \rightarrow J/\psi\pi^+$  yields  $\tau_{B_c^+} = 0.5134 \pm 0.0110 \pm 0.0057$  ps [57]. Each of these is more precise

**Table 2:** Summary of inclusive and exclusive world-average  $b$ -hadron lifetime measurements. For the two  $B_s$  averages, see text below.

Particle	Lifetime [ps]
$B^+$	$1.638 \pm 0.004$
$B^0$	$1.520 \pm 0.004$
$B_s$ (flavor-specific)	$1.511 \pm 0.014$
$B_s$ ( $1/\Gamma_s$ )	$1.510 \pm 0.005$
$B_c^+$	$0.507 \pm 0.009$
$\Lambda_b^0$	$1.466 \pm 0.010$
$\Xi_b^-$	$1.560 \pm 0.040$
$\Xi_b^0$	$1.464 \pm 0.031$
$\Omega_b^-$	$1.57^{+0.23}_{-0.20}$

than the combination of all previous experiments. For precision comparisons with theory, lifetime ratios are more sensitive. Experimentally we find:

$$\frac{\tau_{B^+}}{\tau_{B^0}} = 1.076 \pm 0.004, \quad \frac{\tau_{B_s}}{\tau_{B^0}} = 0.994 \pm 0.004,$$

$$\frac{\tau_{\Lambda_b}}{\tau_{B^0}} = 0.965 \pm 0.007,$$

while theory makes the following predictions [54,58]

$$\frac{\tau_{B^+}}{\tau_{B^0}} = 1.06 \pm 0.02, \quad \frac{\tau_{B_s}}{\tau_{B^0}} = 1.00 \pm 0.01, \quad \frac{\tau_{\Lambda_b}}{\tau_{B^0}} = 0.88 \pm 0.05.$$

The ratio of  $B^+$  to  $B^0$  lifetimes has a precision of better than 1%, and is significantly different from 1.0, in agreement with predictions [54]. The ratio of  $B_s$  to  $B^0$  lifetimes is expected to be very close to 1.0. While early measurements were in mild tension with theory, the high precision measurements using fully reconstructed decays and clear definition of lifetime (see below) are in good agreement with theory [59,60,61]. The  $\Lambda_b$  lifetime has a history of discrepancies. Predictions were higher than data before the introduction of higher-order effects lowered them. The first indication that early measurements of the  $\Lambda_b$  are on low side came from the CDF data [62,63]. The recent measurements from LHC experiments [64,65,66,67] significantly improve precision and favour higher lifetime, much closer to the lifetime of  $B^0$  meson. The most precise measurement of the  $\Lambda_b$  lifetime performed by LHCb uses  $\Lambda_b \rightarrow J/\psi p K^-$  decays and finds  $\tau_{\Lambda_b} = 1.479 \pm 0.009 \pm 0.010$  ps [66]. With new results, the discrepancy between theory and experiment on the  $\Lambda_b$  lifetime can be considered resolved.

Neutral  $B$  mesons are two-component systems similar to neutral kaons, with a light (L) and a heavy (H) mass eigenstate, and independent decay widths  $\Gamma_L$  and  $\Gamma_H$ . The SM predicts a non-zero width difference  $\Delta\Gamma = \Gamma_L - \Gamma_H > 0$  for both  $B_s$  and  $B_d$ . For  $B_d$ ,  $\Delta\Gamma_d/\Gamma_d$  is expected to be  $\sim 0.2\%$ . Analysis of BaBar and DELPHI data on  $CP$ -specific modes of the  $B^0$  yield a combined result:  $\Delta\Gamma_d/\Gamma_d = 0.015 \pm 0.018$  [32]. Recently LHCb determined value of  $\Delta\Gamma_d/\Gamma_d = -0.044 \pm 0.025 \pm 0.011$  [67], which is based on the comparison of lifetimes in the  $B^0 \rightarrow J/\psi K^{*0}(892)$  and  $B^0 \rightarrow J/\psi K_S$  decays. Average including all measurements yields  $\Delta\Gamma_d/\Gamma_d = -0.003 \pm 0.015$ . The

issue is much more interesting for the  $B_s$ , since the SM expectation for  $\Delta\Gamma_s/\Gamma_s$  is of order 10%. This potentially non-negligible difference requires care when defining the  $B_s$  lifetime. As indicated in Table 2, two different lifetimes are defined for the  $B_s$  meson: one is defined as  $1/\Gamma_s$ , where  $\Gamma_s$  is the average width of the two mass eigenstates  $(\Gamma_L + \Gamma_H)/2$ ; the other is obtained from “flavor-specific” (*e.g.*, semileptonic) decays and depends both on  $\Gamma_s$  and  $\Delta\Gamma_s$ . Experimentally, the quantity  $\Delta\Gamma_s$  can be accessed by measuring lifetimes in decays into  $CP$  eigenstates, which in the SM are expected to be close approximations to the mass eigenstates. This has been done with the  $J/\psi\phi$  mode, where the two  $CP$  eigenstates are distinguished by angular distributions, and in  $B_s \rightarrow K^+K^-$  or  $B_s \rightarrow J/\psi f_0(980)$  which are  $CP$ -eigenstates. The current experimental information is dominated by measurements on the  $J/\psi\phi$  mode performed by CDF, D0, ATLAS and LHCb experiments. By appropriately combining all published measurements of  $J/\psi\phi$  lifetimes, flavor-specific lifetimes and effective lifetimes in  $CP$  eigenstates, the HFAG group obtains a world-average  $\Delta\Gamma_s/\Gamma_s = 0.124 \pm 0.011$  [32]; the latest theoretical predictions yield  $\Delta\Gamma_s/\Gamma_s = 0.133 \pm 0.032$  [68], in agreement with measurements within the uncertainties. The constraint from measurements of lifetimes in  $CP$  eigenstates is based on the notion of effective lifetime introduced in Ref. [69]. In this class, measurements in decays  $B_s \rightarrow J/\psi f_0(980)$  [70],  $B_s \rightarrow K^+K^-$  [71] decays are used currently. From the theoretical point of view, the best quantity to use is  $\Delta\Gamma_s/\Delta M_s$ , which is much less affected by hadronic uncertainties [68]. Exploiting the accurate measurement of  $\Delta M_s$  available [72], this can be turned into a SM prediction with an uncertainty of only 20%:  $\Delta\Gamma_s/\Gamma_s = 0.137 \pm 0.027$ . This is likely to be of importance in future comparisons, as the experimental precision improves with the growth of LHC samples. Historically, branching fraction of the decay  $B_s \rightarrow D_s^{(*)+} D_s^{(*)-}$  was used to set an bound on  $\Delta\Gamma_s/\Gamma_s$ , but the method is highly model-dependent and with increased precision of direct determinations it stops to be useful.

The width difference  $\Delta\Gamma_s$  is connected to the  $B_s$  mixing phase  $\phi_s$  by  $\Delta\Gamma_s = \Gamma_{12} \cos \phi_s$ , where  $\Gamma_{12}$  is the off-diagonal element of the decay matrix [6,8,68]. The early measurements by CDF [73] and D0 [74] have produced CL contours in the  $(\phi_s, \Delta\Gamma_s)$  plane, and both observed a mild deviation, in the same direction, from the expectation of the SM of the phase  $\phi_s$  near  $\Delta\Gamma_s = 0$ . The possibility of a large value of  $\phi_s$  has attracted significant interest, as it would be very clean evidence for the existence of new sources of  $CP$  violation beyond the SM. However the latest measurements from CDF [59], D0 [75], ATLAS [76], CMS [77] and LHCb [78], which provide significant improvements over initial measurements, show good agreement with the SM. While most experiments use up to now only  $B_s \rightarrow J/\psi\phi$  decay, LHCb also exploits  $B_s \rightarrow J/\psi\pi^+\pi^-$  decays, which are experimentally determined to be pure  $CP$ -odd and therefore in  $B_s \rightarrow J/\psi\pi^+\pi^-$  decays no angular analysis is needed. It should be noted that in pure  $B_s \rightarrow J/\psi\phi$  decay, there is a two-fold ambiguity in the sign of  $\Delta\Gamma_s$  and  $\phi_s$ . This can be

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resolved using the interference between the decays to  $J/\psi\phi$  and  $J/\psi K^+K^-$ , where  $K^+K^-$  is in relative S-wave state. This has been used by LHCb experiment to determine the sign of  $\Delta\Gamma_s$  to be positive [79] in accordance with SM. The world average value of the  $CP$  violating phase is  $\phi_s = -0.013 \pm 0.037$  [32] without any tension with the SM.

***B* meson decay properties:** Semileptonic  $B$  decays  $B \rightarrow X_c \ell \nu$  and  $B \rightarrow X_u \ell \nu$  provide an excellent way to measure the magnitude of the CKM elements  $|V_{cb}|$  and  $|V_{ub}|$  respectively, because the strong interaction effects are much simplified due to the two leptons in the final state. Both exclusive and inclusive decays can be used with dominant uncertainties being complementary. For exclusive decay analysis, knowledge of the form factors for the exclusive hadronic system  $X_{c(u)}$  is required. For inclusive analysis, it is usually necessary to restrict the available phase-space of the decay products to suppress backgrounds; subsequently uncertainties are introduced in the extrapolation to the full phase-space. Moreover, restriction to a small corner of the phase-space may result in breakdown of the operator-product expansion scheme, thus making theoretical calculations unreliable. One of the recent unexpected results was determination of  $|V_{ub}|$  using  $\Lambda_b^0 \rightarrow p\mu^-\bar{\nu}_\mu$  decays by LHCb [80]. A more detailed discussion of  $B$  semileptonic decays and the extraction of  $|V_{cb}|$  and  $|V_{ub}|$  is given elsewhere in this *Review* [9].

On the other hand, hadronic decays of  $B$  are complicated because of strong interaction effects caused by the surrounding cloud of light quarks and gluons. While this complicates the extraction of CKM matrix elements, it also provides a great opportunity to study perturbative and non-perturbative QCD, hadronization, and Final State Interaction (FSI) effects. Pure-penguin decays were first established by the observation of  $B \rightarrow K^*\gamma$  [81]. Some observed decay modes such as  $B^0 \rightarrow D_s^- K^+$ , may be interpreted as evidence of a  $W$ -exchange process [82]. The evidence for the decay  $B^+ \rightarrow \tau^+\nu$  from Belle [83] and BaBar [84] is the first sign of a pure annihilation decay. There is growing evidence that penguin annihilation processes may be important in decays with two vector mesons in the final state [85].

***Hadronic decays:*** Most of the hadronic  $B$  decays involve  $b \rightarrow c$  transition at the quark level, resulting in a charmed hadron or charmonium in the final state. Other types of hadronic decays are very rare and will be discussed separately in the next section. The experimental results on hadronic  $B$  decays have steadily improved over the past few years, and the measurements have reached sufficient precision to challenge our understanding of the dynamics of these decays. With the good neutral particle detection and hadron identification capabilities of  $B$ -factory detectors, a substantial fraction of hadronic  $B$  decay events can be fully reconstructed. Because of the kinematic constraint of  $\Upsilon(4S)$ , the energy sum of the final-state particles of a  $B$  meson decay is always equal to one half of the total energy in the center of mass frame. As a result, the two variables,  $\Delta E$  (energy difference) and  $M_B$  ( $B$  candidate mass

with a beam-energy constraint) are very effective for suppressing combinatorial background both from  $\Upsilon(4S)$  and  $e^+e^- \rightarrow q\bar{q}$  continuum events. In particular, the energy-constraint in  $M_B$  improves the signal resolution by almost an order of magnitude.

The kinematically clean environment of  $B$  meson decays provides an excellent opportunity to search for new states. For instance, quark-level  $b \rightarrow c\bar{c}s$  decays have been used to search for new charmonium and charm-strange mesons and study their properties in detail. In 2003, BaBar discovered a new narrow charm-strange state  $D_{sJ}^*(2317)$  [86], and CLEO observed a similar state  $D_{sJ}(2460)$  [87]. The properties of these new states were studied in the  $B$  meson decays,  $B \rightarrow DD_{sJ}^*(2317)$  and  $B \rightarrow DD_{sJ}(2460)$  by Belle [88]. Further studies of  $D_{sJ}^{(*)}$  meson production in  $B$  decays have been made by Belle [89] and BaBar [90]. Now these charm-strange meson states are identified as  $D_{s0}^*(2317)$  and  $D_{s1}(2460)$ , respectively.

More recently, Belle observed a new  $D_{sJ}$  meson produced in  $B^+ \rightarrow \bar{D}^0 D_{sJ} \rightarrow \bar{D}^0 D^0 K^+$  [91]. Combined with a subsequent measurement by BaBar [92], the mass and width of this state are determined to be  $2709^{+9}_{-6}$  MeV/ $c^2$  and  $125 \pm 30$  MeV, respectively. An analysis of the helicity angle distribution determines its spin-parity to be  $1^-$ .

A variety of exotic particles have been discovered in  $B$  decays. Belle found the  $X(3872)$  state [93], which is confirmed by CDF [94] and BaBar [95]. Analyzing their full  $\Upsilon(4S)$  data sample, Belle finds a new upper limit on the width of  $X(3872)$  to be  $\Gamma_{X(3872)} < 1.2$  MeV [96], improving on the existing limit by nearly a factor of 2. Radiative decays of  $X(3872)$  can play a crucial role in understanding the nature of the particle. For example, in the molecular model the decay of  $X(3872)$  to  $\psi'\gamma$  is expected to be highly suppressed in comparison to the decay to  $J/\psi\gamma$  [97]. BaBar has seen the evidence for the decay to  $J/\psi\gamma$  [98]. The ratio  $R \equiv \mathcal{B}(X(3872) \rightarrow \psi'\gamma)/\mathcal{B}(X(3872) \rightarrow J/\psi\gamma)$  is measured to be  $3.4 \pm 1.4$  by BaBar [99], while Belle obtains  $R < 2.1$  at 90% CL [100].

Belle has observed a near-threshold enhancement in the  $J/\psi\omega$  invariant mass for  $B \rightarrow J/\psi\omega K$  decays [101]. BaBar has studied  $B \rightarrow J/\psi\pi^+\pi^-K$ , finding an excess of  $J/\psi\pi^+\pi^-$  events with a mass just above 4.2 GeV/ $c^2$ ; this is consistent with the  $Y(4260)$  that was observed by BaBar in ISR (Initial State Radiation) events [102]. A Belle study of  $B \rightarrow \psi'K\pi^\pm$  [103] finds a state called  $X(4430)^\pm$  that decays to  $\psi'\pi^\pm$ . This state was searched for by BaBar with similar sensitivity but was not found [104]. The high statistics study by LHCb experiment confirmed existence of the  $X(4430)^\pm$  in decays  $B \rightarrow \psi'K\pi^\pm$  [105]. Moreover the LHCb experiment demonstrated  $X(4430)^\pm$  resonances character by study of the phase motion and saw evidence for another state. Since it is charged, it could not be a charmonium state. In a Dalitz plot analysis of  $\bar{B}^0 \rightarrow \chi_{c1} K^- \pi^+$ , Belle has observed two resonance-like structures in the  $\chi_{c1}\pi^+$  mass distribution [106], labelled as  $X(4050)^\pm$  and  $X(4250)^\pm$  in this *Review*, while no evidence is found by BaBar in a search with similar sensitivity [107]. Another charge state was seen by Belle in the decay

$B^0 \rightarrow J/\psi K^- \pi^+$  [108]. In the amplitude analysis of the decay  $\Lambda_b^0 \rightarrow J/\psi p K^-$  LHCb experiment also demonstrated that charged states in charmonium region are produced also with non-zero baryon number by observing state decaying to  $J/\psi p$  final state [109].

The hadronic decays  $\bar{B}^0 \rightarrow D^{(*)0} h^0$ , where  $h^0$  stands for light neutral mesons such as  $\pi^0, \eta^{(\prime)}, \rho^0, \omega$ , proceed through color-suppressed diagrams, hence they provide useful tests on the factorization models. Both Belle and BaBar have made comprehensive measurements of such color-suppressed hadronic decays of  $\bar{B}^0$  [110].

Information on  $B_s$  and  $\Lambda_b$  decays is limited, though improving with recent studies of large samples at the Tevatron and LHC experiments. Recent additions are decays of  $B_s \rightarrow J/\psi f_0(980)$  [70,111],  $B_s \rightarrow J/\psi f_2'(1525)$  [112], and  $\Lambda_b \rightarrow \Lambda_c \pi^+ \pi^- \pi^-$  [113]. For the latter, not only the total rate is measured, but also structure involving decays through excited  $\Lambda_c$  and  $\Sigma_c$  baryons.

There have been hundreds of publications on hadronic  $B$  decays to open-charm and charmonium final states mostly from the  $B$ -factory experiments. These results are nicely summarized in a recent report by HFAG [32].

**Rare  $B$  decays:** All  $B$ -meson decays that do not occur through the  $b \rightarrow c$  transition are usually called rare  $B$  decays. These include both semileptonic and hadronic  $b \rightarrow u$  decays that are suppressed at leading order by the small CKM matrix element  $V_{ub}$ , as well as higher-order  $b \rightarrow s(d)$  processes such as electroweak and gluonic penguin decays.

Charmless  $B$  meson decays into two-body hadronic final states such as  $B \rightarrow \pi\pi$  and  $K\pi$  are experimentally clean, and provide good opportunities to probe new physics and search for indirect and direct  $CP$  violations. Since the final state particles in these decays tend to have larger momenta than average  $B$  decay products, the event environment is cleaner than for  $b \rightarrow c$  decays. Branching fractions are typically around  $10^{-5}$ . Over the past decade, many such modes have been observed by BaBar, Belle, and CLEO. More recently, comparable samples of the modes with all charged final particles have been reconstructed in  $p\bar{p}$  collisions by CDF and  $pp$  collisions by LHCb by triggering on the impact parameter of the charged tracks. This has also allowed observation of charmless decays of the  $B_s$ , in final states such as  $\phi\phi$  [114],  $K^+ K^-$  [115], and  $K^- \pi^+$  [116], and of charmless decays of the  $\Lambda_b^0$  baryon [116]. Charmless  $B_s$  modes are related to corresponding  $B^0$  modes by U-spin symmetry, and are determined by similar amplitudes. Combining the observables from  $B_s$  and  $B^0$  modes is a further way of eliminating hadronic uncertainties and extracting relevant CKM information [117].

Because of relatively high-momenta for final state particles, the dominant source of background in  $e^+e^-$  collisions is  $q\bar{q}$  continuum events; sophisticated background suppression techniques exploiting event shape variables are essential for these analyses. In hadron collisions, the dominant background comes

from QCD or partially reconstructed heavy flavors, and is similarly suppressed by a combination of kinematic and isolation requirements. The results are in general consistent among the experiments.

BaBar [118] and Belle [119] have observed the decays  $B^+ \rightarrow \bar{K}^0 K^+$  and  $B^0 \rightarrow K^0 \bar{K}^0$ . The world-average branching fractions are  $\mathcal{B}(B^0 \rightarrow K^0 \bar{K}^0) = (0.96^{+0.20}_{-0.18}) \times 10^{-6}$  and  $\mathcal{B}(B^+ \rightarrow \bar{K}^0 K^+) = (1.36 \pm 0.27) \times 10^{-6}$ . These are the first observations of hadronic  $b \rightarrow d$  transitions, with significance  $> 5\sigma$  for all four measurements.  $CP$  asymmetries have even been measured for these modes, though with large errors.

Most rare decay modes including  $B^0 \rightarrow K^+ \pi^-$  have contributions from both  $b \rightarrow u$  tree and  $b \rightarrow sg$  penguin processes. If the size of the two contributions are comparable, the interference between them may result in direct  $CP$  violation, seen experimentally as a charge asymmetry in the decay rate measurement. BaBar [120], Belle [121], and CDF [115] have measured the direct  $CP$  violating asymmetry in  $B^0 \rightarrow K^+ \pi^-$  decays. The BaBar and Belle measurements constitute observation of direct  $CP$  violation with a significance of more than  $5\sigma$ . The world average for this quantity is now rather precise,  $-0.098 \pm 0.013$ . There are sum rules [122] that relate the decay rates and decay-rate asymmetries between the four  $K\pi$  charge states. The experimental measurements of the other three modes are not yet precise enough to test these sum rules.

There is now evidence for direct  $CP$  violation in three other decays:  $B^+ \rightarrow \rho^0 K^+$  [123],  $B^+ \rightarrow \eta K^+$  [124], and  $B^0 \rightarrow \eta K^{*0}$  [125]. The significance is typically  $3\text{--}4\sigma$ , though the significance for the  $B^+ \rightarrow \eta K^+$  decay is now nearly  $5\sigma$  with the recent Belle measurement [124]. In at least the first two cases, a large direct  $CP$  violation might be expected since the penguin amplitude is suppressed so the tree and penguin amplitudes may have comparable magnitudes.

The decay  $B^0 \rightarrow \pi^+ \pi^-$  can be used to extract the CKM angle  $\alpha$ . This is complicated by the presence of significant contributions from penguin diagrams. An isospin analysis [126] can be used to untangle the penguin complications. The decay  $B^0 \rightarrow \pi^0 \pi^0$ , which is now measured by both BaBar and Belle [127], is crucial in this analysis. Unfortunately the amount of penguin pollution in the  $B \rightarrow \pi\pi$  system is rather large. In the past few years, measurements in the  $B^0 \rightarrow \rho\rho$  system have produced more precise values of  $\alpha$ , since penguin amplitudes are generally smaller for decays with vector mesons. An important ingredient in the analysis is the  $B^0 \rightarrow \rho^0 \rho^0$  branching fraction. The average of measurements from BaBar and Belle [128] yields a branching fraction of  $(0.73 \pm 0.28) \times 10^{-6}$ . This is only 3% of the  $\rho^+ \rho^-$  branching fraction, much smaller than the corresponding ratio ( $\sim 10\%$ ) in the  $\pi\pi$  system.

The decay  $B \rightarrow a_1 \pi$  has been seen by BaBar. An analysis of the time evolution of this decay [129] together with measurements of other related decays has been used to measure the CKM angle  $\alpha$  [130] in agreement with the more precise measurements from the  $\rho\rho$  system.

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Since  $B \rightarrow \rho\rho$  has two vector mesons in the final state, the  $CP$  eigenvalue of the final state depends on the longitudinal polarization fraction  $f_L$  for the decay. Therefore, a measurement of  $f_L$  is needed to extract the CKM angle  $\alpha$ . Both BaBar and Belle have measured  $f_L$  for the decays  $\rho^+\rho^-$  [131] and  $\rho^+\rho^0$  [132] and in both cases the measurements show  $f_L > 0.9$ , making a complete angular analysis unnecessary.

By analyzing the angular distributions of the  $B$  decays to two vector mesons, we can learn a lot about both weak- and strong-interaction dynamics in  $B$  decays. Decays that are penguin-dominated surprisingly have values of  $f_L$  near 0.5. The list of such decays has now grown to include  $B \rightarrow \phi K^*$ ,  $B \rightarrow \rho K^*$ , and  $B \rightarrow \omega K^*$ . The reasons for this "polarization puzzle" are not fully understood. A detailed description of the angular analysis of  $B$  decays to two vector mesons can be found in a separate mini-review [133] in this *Review*.

There has been substantial progress in measurements of many other rare- $B$  decays. The decay  $B \rightarrow \eta' K$  stood out as the largest rare- $B$  decay for many years. The reasons for the large rate are now largely understood [14,134]. However, there are now measurements of several 3-body or quasi-3-body modes with similarly large branching fractions. States seen so far include  $K\pi\pi$  (three charge states) [135],  $KKK$  (four charge states) [136], and  $K^*\pi\pi$  (two charged states) [137]. Many of these analyses now include Dalitz plot treatments with many intermediate resonances. There has also been an observation of the decay  $B^+ \rightarrow K^+ K^- \pi^+$  by BaBar [138], noteworthy because an even number of kaons is typically indicative of suppressed  $b \rightarrow d$  transitions as discussed above.

Belle [83] and BaBar [84] have found evidence for  $B^+ \rightarrow \tau^+ \nu$ ; the average branching fraction is  $(1.14 \pm 0.27) \times 10^{-4}$ . This is somewhat larger than, though consistent with, the value expected in the SM. This is the first observation of a pure annihilation decay. A substantial region of parameter space of charged Higgs mass vs.  $\tan\beta$  is excluded by the measurements of this mode.

**Electroweak penguin decays:** More than 20 years have passed since the CLEO experiment first observed an exclusive radiative  $b \rightarrow s\gamma$  transition,  $B \rightarrow K^*(892)\gamma$  [81], thus providing the first evidence for the one-loop FCNC electromagnetic penguin decay. Using much larger data samples, both Belle and BaBar have updated this analysis [139] with an average branching fraction  $\mathcal{B}(B^0 \rightarrow K^{*0}\gamma) = (43.3 \pm 1.5) \times 10^{-6}$ , and have added several new decay modes such as  $B \rightarrow K_1\gamma$ ,  $K_2^*(1430)\gamma$ , *etc.* [140]. With a sample of 24 fb $^{-1}$  at  $\Upsilon(5S)$ , Belle observed the radiative penguin decay of  $B_s \rightarrow \phi\gamma$  [141]. The decay  $B_s \rightarrow \phi\gamma$  was also seen at LHCb with higher statistics [142]. The two measurements give average branching fraction of  $(36 \pm 4) \times 10^{-6}$ .

Compared to  $b \rightarrow s\gamma$ , the  $b \rightarrow d\gamma$  transitions such as  $B \rightarrow \rho\gamma$ , are suppressed by the small CKM element  $V_{td}$ . Both Belle and BaBar have observed these decays [17,18]. The world average  $\mathcal{B}(B \rightarrow (\rho, \omega)\gamma) = (1.30 \pm 0.23) \times 10^{-6}$ . This can be

used to calculate  $|V_{td}/V_{ts}|$  [143]; the measured values are  $0.233^{+0.033}_{-0.032}$  from BaBar [18] and  $0.195^{+0.025}_{-0.024}$  from Belle [17].

The observed radiative penguin branching fractions can constrain a large class of SM extensions [144]. However, due to the uncertainties in the hadronization, only the inclusive  $b \rightarrow s\gamma$  rate can be reliably compared with theoretical calculations. This rate can be measured from the endpoint of the inclusive photon spectrum in  $B$  decay. By combining the measurements of  $B \rightarrow X_s\gamma$  from CLEO, BaBar, and Belle experiments [145,146,147], HFAG obtains the new average:  $\mathcal{B}(B \rightarrow X_s\gamma) = (3.43 \pm 0.21 \pm 0.07) \times 10^{-4}$  [32] for  $E_\gamma \geq 1.6$  GeV, which averages over  $B^+$  and  $B^0$ . Consistent but less precise results have been reported by ALEPH for inclusive  $b$ -hadrons produced at the  $Z$ , which includes also small fraction of  $B_s$  and  $\Lambda_b$  hadrons. The measured branching fraction can be compared to theoretical calculations. Recent calculations of  $\mathcal{B}(b \rightarrow s\gamma)$  at NNLO level predict the values of  $(3.15 \pm 0.23) \times 10^{-4}$  [148] and  $(2.98 \pm 0.26) \times 10^{-4}$  [149], where the latter is calculated requiring  $E_\gamma \geq 1.6$  GeV.

The  $CP$  asymmetry in  $b \rightarrow s\gamma$  is extensively studied theoretically both in the SM and beyond [150]. According to the SM, the  $CP$  asymmetry in  $b \rightarrow s\gamma$  is smaller than 1%, but some non-SM models allow significantly larger  $CP$  asymmetry ( $\sim 10\%$ ) without altering the inclusive branching fraction. The current world average is  $A_{CP} = -0.008 \pm 0.029$ , again dominated by BaBar and Belle [151,146]. In addition to the  $CP$  asymmetry, BaBar also measured the isospin asymmetry  $\Delta_{0-} = -0.01 \pm 0.06$  in  $b \rightarrow s\gamma$  measured using sum of exclusive decays [152]. Alternative measurement using full reconstruction of the companion  $B$  in the hadronic decay modes yields consistent, but less precise result [153].

In addition, all three experiments have measured the inclusive photon energy spectrum for  $b \rightarrow s\gamma$ , and by analyzing the shape of the spectrum they obtain the first and second moments for photon energies. Belle has measured these moments covering the widest range in the photon energy ( $1.7 < E_\gamma < 2.8$  GeV) [147]. The measurement by BaBar has slightly smaller range with lower limit at 1.8 GeV [154]. These results can be used to extract non-perturbative HQET parameters that are needed for precise determination of the CKM matrix element  $V_{ub}$ .

Additional information on FCNC processes can be obtained from  $b \rightarrow s\ell^+\ell^-$  decays, which are mediated by electroweak penguin and  $W$ -box diagrams. Measurements at Belle and BaBar suffered from low statistics and therefore they typically provide average between charged and neutral  $B$  mesons as well as between  $e^+e^-$  and  $\mu^+\mu^-$  final states [155,156]. The total branching fraction measured at B-factories for  $B \rightarrow K\ell^+\ell^-$  is  $(0.45 \pm 0.04) \times 10^{-6}$  and for  $B \rightarrow K^*(892)\ell^+\ell^-$  is  $(1.05 \pm 0.10) \times 10^{-6}$ . Measurements at B-factories were complemented by CDF [157], which used only muons in the final state. While precision at CDF was similar to B-factories, it had access also to  $B_s \rightarrow \phi\mu^+\mu^-$  and  $\Lambda_b \rightarrow \Lambda\mu^+\mu^-$  decays, which were observed for the first time [157,158] and confirmed by LHCb [159,160,161].  $B$ -factory experiments also measured



the branching fractions for inclusive  $B \rightarrow X_s \ell^+ \ell^-$  decays [162], with an average of  $(3.66^{+0.76}_{-0.77}) \times 10^{-6}$  [163]. In  $b \rightarrow s \ell^+ \ell^-$  decays, the angular analysis provides several interesting observables, which can be studied as function of dilepton invariant mass squared,  $q^2$ . While first measurements were done by Belle, Babar and CDF, real advance of these measurements came with LHC experiments, where samples available are significantly larger than before. The best known of angular observables is forward-backward asymmetry, which was measured in  $B \rightarrow K^*(892) \ell^+ \ell^-$  by several experiments having access to the decay [155,164,165,166,167,168] with most precise measurement coming from LHCb [169]. Measurements of the  $CP$  asymmetries [156,170,171], the isospin asymmetry [155,156,172] and several other angular observables [167,173] are possible in this class of decays. While most of the measurements agree with the SM, the differential branching fractions, isospin asymmetry in  $B \rightarrow K \mu^+ \mu^-$  and the other angular observable  $P'_5$  [174] measured by the LHCb exhibit small tension with the SM expectation. Although initially the angular analyses were mainly concentrating on the decay  $B \rightarrow K^*(892) \ell^+ \ell^-$ , with large-statistics samples available at LHC, angular analyses of the  $B \rightarrow K \mu^+ \mu^-$  [175],  $B_s \rightarrow \phi \mu^+ \mu^-$  [159] and  $\Lambda_b^0 \rightarrow \Lambda \mu^+ \mu^-$  [161] decays were also performed, with the results being consistent with the SM.

With the data samples available at LHC, lepton universality in  $b \rightarrow s \ell^+ \ell^-$  can be tested. While in the standard model decays to electron-positron and muon pairs are expected to be same up to small corrections due to different masses of leptons, in extensions of the SM this does not have to hold. With this aim the angular analysis of  $B^0 \rightarrow K^{*0} e^+ e^-$  decays was performed by LHCb at low dilepton invariant masses [176]. Most notable result on lepton universality test is the ratio of branching fractions between  $B^+ \rightarrow K^+ \mu^+ \mu^-$  and  $B^+ \rightarrow K^+ e^+ e^-$  at LHCb, which shows  $2.6\sigma$  discrepancy with the SM [177].

Finally the decays  $B_{(s)}^0 \rightarrow e^+ e^-$  and  $\mu^+ \mu^-$  are interesting since they only proceed at second order in weak interactions in the SM, but may have large contributions from supersymmetric loops, proportional to  $(\tan \beta)^6$ . First limits were published 30 years ago and since then experiments at Tevatron,  $B$ -factories and LHC gradually improved those and effectively excluded whole models of new physics and significantly constrained allowed parameter space of others. For the decays to  $\mu^+ \mu^-$ , Tevatron experiments pushed the limits down to roughly factor of 5-10 above the SM expectation [178,179]. The long journey in the search for these decays culminated in 2012, when first evidence for  $B_s \rightarrow \mu^+ \mu^-$  decay was seen [180]. Currently LHCb [181] and CMS [182] observe this decay with significance between 4 and 5 standard deviations. The measured branching fraction is  $(2.9^{+1.1}_{-1.0}) \times 10^{-9}$  at LHCb and  $(3.0^{+1.0}_{-0.9}) \times 10^{-9}$  at CMS, both in agreement with the SM expectation. The combination of the data from CMS and LHCb yields branching fraction  $(2.8^{+0.7}_{-0.6}) \times 10^{-9}$  for  $B_s \rightarrow \mu^+ \mu^-$  decay [183]. The statistical significance of the combined signal is  $6.2\sigma$ . For the  $B^0 \rightarrow \mu^+ \mu^-$  decay, combined analysis of CMS and LHCb

data gives branching fraction  $(3.9^{+1.6}_{-1.4}) \times 10^{-10}$  and statistical significance of  $3.2\sigma$ . The measured branching fraction for  $B_s$  is compatible at  $1.2\sigma$  with the SM. For the  $B^0$ , the measurement is about  $2.2\sigma$  above the SM prediction. Recently ATLAS [184] reported a study of  $B^0 \rightarrow \mu^+ \mu^-$  and  $B_s^0 \rightarrow \mu^+ \mu^-$  decays. For  $B^0$  an upper limit on  $B(B^0 \rightarrow \mu^+ \mu^-) < 4.2 \times 10^{-10}$  is set at 95% C.L. For  $B_s^0$  the result is  $B(B_s^0 \rightarrow \mu^+ \mu^-) = (0.9^{+1.1}_{-0.8}) \times 10^{-9}$ . An upper limit  $B(B_s^0 \rightarrow \mu^+ \mu^-) < 3.0 \times 10^{-9}$  at 95% CL, lower than the SM prediction, and in agreement with the measurement of CMS and LHCb. The limits for the  $e^+ e^-$  modes are:  $< 2.8 \times 10^{-7}$  and  $< 8.3 \times 10^{-8}$ , respectively, for  $B_s$  and  $B^0$  [185]. The searches were also performed for lepton flavour violating decays to two leptons with best limits in  $e^\pm \mu^\mp$  channel, where limits are  $< 3.7 \times 10^{-9}$  for  $B^0$  and  $< 1.4 \times 10^{-8}$  for  $B_s$ , at 95% confidence level [186].

**Summary and Outlook:** The study of  $B$  mesons continues to be one of the most productive fields in particle physics. With the two asymmetric  $B$ -factory experiments Belle and BaBar, we now have a combined data sample of well over  $1 \text{ ab}^{-1}$ .  $CP$  violation has been firmly established in many decays of  $B$  mesons. Evidence for direct  $CP$  violation has been observed. Many rare decays resulting from hadronic  $b \rightarrow u$  transitions and  $b \rightarrow s(d)$  penguin decays have been observed, and the emerging pattern is still full of surprises. Despite the remarkable successes of the  $B$ -factory experiments, many fundamental questions in the flavor sector remain unanswered.

At Fermilab, CDF and D0 each has accumulated about  $10 \text{ fb}^{-1}$ , which is the equivalent of about  $10^{12}$   $b$ -hadrons produced. In spite of the low trigger efficiency of hadronic experiments, a selection of modes have been reconstructed in large quantities, giving a start to a program of studies on  $B_s$  and  $b$ -flavored baryons, in which a first major step has been the determination of the  $B_s$  oscillation frequency.

As Tevatron and  $B$ -factories stop their taking data, the new experiments at the LHC have become very active. The LHC accelerator performed very well in 2011 and 2012. The general purpose experiments ATLAS and CMS collected about  $25 \text{ fb}^{-1}$  while LHCb collected about  $3 \text{ fb}^{-1}$ . After two years of consolidation, the LHC restarted in 2015 and experiments expect to double their  $b$ -hadrons samples during 2016. LHCb, which is almost fully dedicated to the studies of  $b$ - and  $c$ -hadrons, has a data sample that is for many decays larger than the sum of all previous experiments. Of particular note is that after many years of search for the decay  $B_s \rightarrow \mu^+ \mu^-$  the LHCb and CMS experiments finally observed this decay in the combined analysis of their data [183].

In addition, the preparation of the next generation high-luminosity  $B$ -factory at KEK is in its final stages with first physics data taking expected in 2017. The aim to increase sample to  $\sim 50 \text{ ab}^{-1}$  will make it possible to explore the indirect evidence of new physics beyond the SM in the heavy-flavor particles ( $b$ ,  $c$ , and  $\tau$ ), in a way that is complementary to the LHC. In the same time, LHCb Collaboration is working

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on the upgrade of its detector, which should be installed in 2018 and 2019. Aim of the upgrade is to increase flexibility of trigger, which will allow to significantly increase instantaneous luminosity and possibly integrate about  $50 \text{ fb}^{-1}$  of data.

These experiments promise a rich spectrum of rare and precise measurements that have the potential to fundamentally affect our understanding of the SM and *CP*-violating phenomena.

### References

1. M. Kobayashi and T. Maskawa, *Prog. Theor. Phys.* **49**, 652 (1973).
2. S. W. Herb *et al.*, *Phys. Rev. Lett.* **39**, 252 (1977).
3. B. Aubert *et al.* (BaBar Collab.), *Phys. Rev. Lett.* **87**, 091801 (2001).
4. K. Abe *et al.* (Belle Collab.), *Phys. Rev. Lett.* **87**, 091802 (2001).
5. Currently two different notations ( $\phi_1, \phi_2, \phi_3$ ) and ( $\alpha, \beta, \gamma$ ) are used in the literature for CKM unitarity angles. In this mini-review, we use the latter notation following the other mini-reviews in this *Review*. The two notations are related by  $\phi_1 = \beta$ ,  $\phi_2 = \alpha$  and  $\phi_3 = \gamma$ .
6. See the “*CP* Violation in Meson Decays” by D. Kirkby and Y. Nir in this *Review*.
7. See the “CKM Quark Mixing Matrix,” by A. Cecucci, Z. Ligeti, and Y. Sakai, in this *Review*.
8. See the “Review on *B*- $\bar{B}$  Mixing,” by O. Schneider in this *Review*.
9. See the “Determination of  $|V_{cb}|$  and  $|V_{ub}|$ ,” by R. Kowalewski and T. Mannel in this *Review*.
10. The  $B_c$  is a special case, where a weak decay of the *c* quark is also possible, but the spectator model still applies.
11. B. Grinstein, *Nucl. Phys.* **B339**, 253 (1990); H. Georgi, *Phys. Lett.* **B240**, 447 (1990); A.F. Falk *et al.*, *Nucl. Phys.* **B343**, 1 (1990); E. Eichten and B. Hill, *Phys. Lett.* **B234**, 511 (1990).
12. “Heavy-Quark and Soft-Collinear Effective Theory” by C.W. Bauer and M. Neubert in this *Review*.
13. M. Neubert, “Aspects of QCD Factorization,” *hep-ph/0110093*, *Proceedings of HF9*, Pasadena (2001) and references therein; Z. Ligeti *et al.*, *Phys. Lett.* **B507**, 142 (2001).
14. M. Beneke *et al.*, *Phys. Rev. Lett.* **83**, 1914 (1999); *Nucl. Phys.* **B591**, 313 (2000); *Nucl. Phys.* **B606**, 245 (2001); M. Beneke and M. Neubert, *Nucl. Phys.* **B675**, 333 (2003).
15. Y.Y. Keum, H.-n. Li, and A.I. Sanda, *Phys. Lett.* **B504**, 6 (2001); *Phys. Rev.* **D63**, 054008 (2001); Y.Y. Keum and H.-n. Li, *Phys. Rev.* **D63**, 074006 (2001); C.D. Lü, K. Ukai, and M.Z. Yang, *Phys. Rev.* **D63**, 074009 (2001); C.D. Lü and M.Z. Yang, *Eur. Phys. J.* **C23**, 275 (2002).
16. C.W. Bauer, S. Fleming, and M.E. Luke, *Phys. Rev.* **D63**, 014006 (2001); C.W. Bauer *et al.*, *Phys. Rev.* **D63**, 114020 (2001); C.W. Bauer and I.W. Stewart, *Phys. Lett.* **B516**, 134 (2001).
17. N. Taniguchi *et al.* (Belle Collab.), *Phys. Rev. Lett.* **101**, 111801 (2008).
18. B. Aubert *et al.* (BaBar Collab.), *Phys. Rev.* **D78**, 112001 (2008).
19. F. Abe *et al.* (CDF Collab.), *Phys. Rev. Lett.* **81**, 2432 (1998); F. Abe *et al.* (CDF Collab.), *Phys. Rev.* **D58**, 112004 (1998).
20. D. Acosta *et al.* (CDF Collab.), *Phys. Rev. Lett.* **96**, 082002 (2006).
21. R. Aaij *et al.* (LHCb Collab.), *Phys. Rev.* **D87**, 112012 (2013).
22. B. Aubert *et al.* (BABAR Collab.), *Phys. Rev. Lett.* **101**, 071801 (2008) [Erratum-ibid. **102**, 029901 (2009)].
23. G. Bonvicini, *et al.* (CLEO Collab.), *Phys. Rev.* **D81**, 031104 (2010).
24. See the note on “Naming scheme for hadrons,” by M. Roos and C.G. Wohl in this *Review*.
25. A. Abulencia *et al.* (CDF Collab.), *Phys. Rev.* **D75**, 012010 (2007), and references therein; R. Aaij *et al.* (LHCb Collab.), *JHEP* **1308**, 117 (2013); S. Chatrchyan *et al.* (CMS Collab.), *JHEP* **1206**, 110 (2012), and references therein; G. Aad *et al.* (ATLAS Collab.), *JHEP* **1310**, 042 (2013), and references therein; B. Abelev *et al.* (ALICE Collab.), *Phys. Lett.* **B721**, 13 (2013).
26. M. Cacciari *et al.*, *JHEP* **9805**, 007 (1998); S. Frixione and B. R. Webber, *JHEP* **0206**, 029 (2002); M. Cacciari *et al.*, *JHEP* **0407**, 033 (2004); M. Cacciari *et al.*, *JHEP* **0604**, 006 (2006), and references therein; M. Cacciari *et al.*, *JHEP* **1210**, 137 (2012).
27. B. Barish *et al.* (CLEO Collab.), *Phys. Rev. Lett.* **76**, 1570 (1996).
28. B. Aubert *et al.* (BaBar Collab.), *Phys. Rev. Lett.* **96**, 232001 (2006); A. Sokolov *et al.* (Belle Collab.), *Phys. Rev.* **D75**, 071103 (R) (2007).
29. J.P. Alexander *et al.* (CLEO Collab.), *Phys. Rev. Lett.* **86**, 2737 (2001).
30. N.C. Hastings *et al.* (Belle Collab.), *Phys. Rev.* **D67**, 052004 (2003).
31. B. Aubert *et al.* (BaBar Collab.), *Phys. Rev. Lett.* **95**, 042001 (2005).
32. Y. Amhis *et al.* (Heavy Flavor Averaging Group), [arXiv:1412.7515](https://arxiv.org/abs/1412.7515), and online update at <http://www.slac.stanford.edu/xorg/hfag/>.
33. G.S. Huang *et al.* (CLEO Collab.), *Phys. Rev.* **D75**, 012002 (2007).
34. R. Louvot *et al.* (Belle Collab.), *Phys. Rev. Lett.* **102**, 021801 (2009).
35. M. Lusignoli, M. Masetti, and S. Petrarca, *Phys. Lett.* **B266**, 142 (1991); K. Cheung, *Phys. Lett.* **B472**, 408 (2000).
36. T. Aaltonen *et al.* (CDF Collab.), *Phys. Rev.* **D77**, 072003 (2008).
37. R. Aaij *et al.* (LHCb Collab.), *Phys. Rev.* **D85**, 032008 (2012).
38. P. Abreu *et al.* (DELPHI Collab.), *Phys. Lett.* **B345**, 598 (1995).
39. T. Aaltonen *et al.* (CDF Collab.), *Phys. Rev. Lett.* **102**, 102003 (2009); V.M. Abazov *et al.* (D0 Collab.), *Phys. Rev. Lett.* **99**, 172001 (2007).
40. R. Akers *et al.* (OPAL Collab.), *Z. Phys.* **C66**, 19 (1995).
41. T. Aaltonen *et al.* (CDF Collab.), *Phys. Rev. Lett.* **100**, 082001 (2008); V.M. Abazov *et al.* (D0 Collab.), *Phys. Rev. Lett.* **100**, 082002 (2008).

42. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **110**, 151803 (2013).
43. D. Buskulic *et al.* (ALEPH Collab.), Phys. Lett. **B384**, 449 (1996); P. Abreu *et al.* (DELPHI Collab.), Z. Phys. **C68**, 541 (1995).
44. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **99**, 202001 (2007); T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D85**, 092011 (2012).
45. V.M. Abazov *et al.* (D0 Collab.), Phys. Rev. Lett. **99**, 052001 (2007).
46. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **99**, 052002 (2007).
47. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **107**, 102001 (2011).
48. V. M. Abazov *et al.* (D0 Collab.), Phys. Rev. Lett. **101**, 232002 (2008).
49. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. D **80**, 072003 (2009).
50. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **110**, 182001 (2013).
51. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **108**, 252002 (2012).
52. V. M. Abazov *et al.* (D0 Collab.), arXiv:1602.07588, submitted to Phys. Rev. Lett..
53. The LHCb Collaboration, LHCb-CONF-2016-004, CERN-LHCb-CONF-2016-004..
54. C. Tarantino, Eur. Phys. J. **C33**, S895 (2004); F. Gabbiani *et al.*, Phys. Rev. **D68**, 114006 (2003); F. Gabbiani *et al.*, Phys. Rev. **D70**, 094031 (2004).
55. C.H. Chang *et al.*, Phys. Rev. **D64**, 014003 (2001); V.V. Kiselev, A.E. Kovalsky, and A.K. Likhoded, Nucl. Phys. **B585**, 353 (2000); V.V. Kiselev, arXiv:hep-ph/0308214, and references therein.
56. R. Aaij *et al.* (LHCb Collab.), Eur. Phys. J. **C74**, 2839 (2014).
57. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B742**, 29 (2015).
58. I.I. Bigi *et al.*, in *B Decays*, 2nd ed., S. Stone (ed.), World Scientific, Singapore, 1994.
59. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **109**, 171802 (2012).
60. G. Aad *et al.* (ATLAS Collab.), JHEP **12**, 072 (2012).
61. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. **D87**, 112010 (2013).
62. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **104**, 102002 (2010).
63. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **106**, 121804 (2011).
64. S. Chatrchyan *et al.* (CMS Collab.), JHEP **07**, 163 (2013).
65. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D87**, 032002 (2013).
66. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B734**, 122 (2014).
67. R. Aaij *et al.* (LHCb Collab.), JHEP **1404**, 114 (2014).
68. A. Lenz and U. Nierste, JHEP **0706**, 072 (2007) and numerical update in arXiv:1102.4274 [hep-ph].
69. R. Fleischer and R. Knegjens, Eur. Phys. J. **C71**, 1789 (2011).
70. A. Abulencia *et al.* (CDF Collab.), Phys. Rev. **D84**, 052012 (2012).
71. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B707**, 349 (2012).
72. A. Abulencia *et al.* (CDF Collab.), Phys. Rev. Lett. **97**, 242003 (2006).
73. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **100**, 121803 (2008).
74. V.M. Abazov *et al.* (D0 Collab.), Phys. Rev. **D76**, 057101 (2007).
75. V. M. Abazov *et al.* (D0 Collab.), Phys. Rev. **D85**, 032006 (2012).
76. G. Aad *et al.* (ATLAS Collab.), Phys. Rev. **D90**, 052007 (2014).
77. V. Khachatryan *et al.* (CMS Collab.), Phys. Lett. **B757**, 97 (2016).
78. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **114**, 041801 (2015).
79. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **108**, 241801 (2012).
80. R. Aaij *et al.* (LHCb Collab.), Nature Phys. **11**, 743 (2015).
81. R. Ammar *et al.* (CLEO Collab.), Phys. Rev. Lett. **71**, 674 (1993).
82. P. Krokovny *et al.* (Belle Collab.), Phys. Rev. Lett. **89**, 231804 (2002); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **98**, 081801 (2007).
83. B. Kronenbitter *et al.* (Belle Collab.), Phys. Rev. **D92**, 051102 (2015); I. Adachi *et al.* (Belle Collab.), Phys. Rev. Lett. **110**, 131801 (2013).
84. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D88**, 031102 (2013); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D81**, 051101 (2010).
85. M. Beneke, J. Rohrer, and D. Yang, Nucl. Phys. **B774**, 64 (2007).
86. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **90**, 242001 (2003).
87. D. Besson *et al.* (CLEO Collab.), Phys. Rev. **D68**, 032002 (2003).
88. P. Krokovny *et al.* (Belle Collab.), Phys. Rev. Lett. **91**, 262002 (2003).
89. Y. Mikami *et al.* (Belle Collab.), Phys. Rev. Lett. **92**, 012002 (2004).
90. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **93**, 181801 (2004).
91. J. Brodzicka *et al.* (Belle Collab.), Phys. Rev. Lett. **100**, 092001 (2008).
92. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D80**, 092003 (2009).
93. S.-K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **91**, 262001 (2003).
94. D. Acosta *et al.* (CDF II Collab.), Phys. Rev. Lett. **93**, 072001 (2004).
95. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D71**, 071103 (2005).
96. S.-K. Choi *et al.* (Belle Collab.), Phys. Rev. **D84**, 052004 (2011).
97. E.S. Swanson, Phys. Rep. **429**, 243 (2006).
98. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D74**, 071101 (2006).

# Meson Particle Listings

## *b*-flavored hadrons

99. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **102**, 132001 (2009).
100. V. Bhardwaj *et al.* (Belle Collab.), Phys. Rev. Lett. **107**, 091803 (2011).
101. S.-K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **94**, 182002 (2005).
102. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **95**, 142001 (2005).
103. S.-K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **100**, 142001 (2008); R. Mizuk *et al.* (Belle Collab.), Phys. Rev. **D80**, 031104 (2009).
104. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D79**, 112001 (2009).
105. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **112**, 222002 (2014); R. Aaij *et al.* (LHCb Collab.), Phys. Rev. **D92**, 112009 (2015).
106. R. Mizuk *et al.* (Belle Collab.), Phys. Rev. **D78**, 072004 (2008).
107. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D85**, 052003 (2012).
108. K. Chilikin *et al.* (Belle Collab.), Phys. Rev. **D90**, 112009 (2014).
109. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **115**, 072001 (2015).
110. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D84**, 112007 (2011); S. Blyth *et al.* (Belle Collab.), Phys. Rev. **D74**, 092002 (2006).
111. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B698**, 115 (2011); J. Li *et al.* (Belle Collab.), Phys. Rev. Lett. **106**, 121802 (2011); V. M. Abazov *et al.* (D0 Collab.), Phys. Rev. **D85**, 011103 (2012).
112. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **707**, 497 (2012).
113. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. **D84**, 092001 (2011); *ibid.* Phys. Rev. **D85**, 039904 (2012); T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D85**, 032003 (2012).
114. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **107**, 261802 (2011).
115. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **106**, 181802 (2011).
116. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **103**, 031801 (2009).
117. R. Fleischer, Phys. Lett. **B459**, 306 (1999); D. London and J. Matias, Phys. Rev. **D70**, 031502 (2004).
118. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **97**, 171805 (2006).
119. S.-W. Lin *et al.* (Belle Collab.), Phys. Rev. Lett. **98**, 181804 (2007).
120. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **99**, 021603 (2007).
121. S.-W. Lin *et al.* (Belle Collab.), Nature **452** 332(2008).
122. See for example M. Gronau and J.L. Rosner, Phys. Rev. **D71**, 074019 (2005); M. Gronau, Phys. Lett. **B627**, 82 (2005).
123. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D78**, 012004 (2008); A. Garmash *et al.* (Belle Collab.), Phys. Rev. Lett. **96**, 251803 (2006).
124. C.-T. Hoi *et al.* (Belle Collab.), Phys. Rev. Lett. **108**, 031801 (2012); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D80**, 112002 (2009).
125. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **97**, 201802 (2006); C.H. Wang *et al.* (Belle Collab.), Phys. Rev. **D75**, 092005 (2007).
126. M. Gronau and D. London, Phys. Rev. Lett. **65**, 3381 (1990).
127. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D87**, 052009 (2013); Y. Chao *et al.* (Belle Collab.), Phys. Rev. Lett. **94**, 181803 (2008).
128. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D78**, 071104 (2008); C.C. Chiang *et al.* (Belle Collab.), Phys. Rev. **D78**, 111102 (2008).
129. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **98**, 181803 (2007).
130. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D81**, 052009 (2010).
131. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D76**, 052007 (2007); A. Somov *et al.* (Belle Collab.), Phys. Rev. Lett. **96**, 171801 (2006).
132. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **102**, 141802 (2009); J. Zhang *et al.* (Belle Collab.), Phys. Rev. Lett. **91**, 221801 (2003).
133. See the “Polarization in *B* Decays,” by A. Gritsan and J.G. Smith in this *Review*.
134. A. Williamson and J. Zupan, Phys. Rev. **D74**, 014003 (2006).
135. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D78**, 012004 (2008); A. Garmash *et al.* (Belle Collab.), Phys. Rev. Lett. **96**, 251803 (2006); P. Chang *et al.* (Belle Collab.), Phys. Lett. **B599**, 148 (2004); J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D83**, 112010 (2011); A. Garmash *et al.* (Belle Collab.), Phys. Rev. **D75**, 012006 (2007); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D80**, 112001 (2009).
136. A. Garmash *et al.* (Belle Collab.), Phys. Rev. **D71**, 092003 (2005); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D74**, 032003 (2006); A. Garmash *et al.* (Belle Collab.), Phys. Rev. **D69**, 012001 (2004); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **93**, 181805 (2004); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **95**, 011801 (2005).
137. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D74**, 051104R (2006); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D76**, 071104R (2007).
138. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **99**, 221801 (2007).
139. M. Nakao *et al.* (Belle Collab.), Phys. Rev. **D69**, 112001 (2004); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **103**, 211802 (2009).
140. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D70**, 091105R (2004); H. Yang *et al.* (Belle Collab.), Phys. Rev. Lett. **94**, 111802 (2005); S. Nishida *et al.* (Belle Collab.), Phys. Lett. **B610**, 23 (2005); B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D74**, 031102R (2004).
141. J. Wicht *et al.* (Belle Collab.), Phys. Rev. Lett. **100**, 121801 (2008).
142. R. Aaij *et al.* (LHCb Collab.), Nucl. Phys. **B867**, 1 (2013).
143. A. Ali *et al.*, Phys. Lett. **B595**, 323 (2004); P. Ball, G. Jones, and R. Zwicky, Phys. Rev. **D75**, 054004 (2007).
144. J.L. Hewett, Phys. Rev. Lett. **70**, 1045 (1993).

- 
145. S. Chen *et al.* (CLEO Collab.), Phys. Rev. Lett. **87**, 251807 (2001).
146. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D86**, 112008 (2012).
147. A. Limosani *et al.* (Belle Collab.), Phys. Rev. Lett. **103**, 241801 (2009).
148. M. Misiak *et al.*, Phys. Rev. Lett. **98**, 022002 (2007).
149. T. Becher and M. Neubert, Phys. Rev. Lett. **98**, 022003 (2007).
150. L. Wolfenstein and Y.L. Wu, Phys. Rev. Lett. **73**, 2809 (1994); H.M. Asatrian and A. Ioannisian, Phys. Rev. **D54**, 5642 (1996); M. Ciuchini *et al.*, Phys. Lett. **B388**, 353 (1996); S. Baek and P. Ko, Phys. Rev. Lett. **83**, 488 (1998); A.L. Kagan and M. Neubert, Phys. Rev. **D58**, 094012 (1998); K. Kiers *et al.*, Phys. Rev. **D62**, 116004 (2000).
151. S. Nishida *et al.* (Belle Collab.), Phys. Rev. Lett. **93**, 031803 (2004).
152. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D72**, 052004 (2005).
153. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D77**, 051103 (2008).
154. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. Lett. **109**, 191801 (2012).
155. J.-T. Wei *et al.* (Belle Collab.), Phys. Rev. Lett. **103**, 171801 (2009).
156. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D86**, 032012 (2012).
157. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **107**, 201802 (2011).
158. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **106**, 161801 (2011).
159. R. Aaij *et al.* (LHCb Collab.), JHEP **1307**, 084 (2013); R. Aaij *et al.* (LHCb Collab.), JHEP **1509**, 179 (2015).
160. R. Aaij *et al.* (LHCb Collab.), Phys. Lett. **B725**, 25 (2013).
161. R. Aaij *et al.* (LHCb Collab.), JHEP **1506**, 115 (2015).
162. Y. Sato *et al.* (Belle Collab.), Phys. Rev. **D93**, 032008 (2016); J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. Lett. **112**, 211802 (2014).
163. The average is calculated by HFAG [32] including the recent unpublished value by Belle.
164. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **108**, 081807 (2012).
165. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **102**, 091803 (2009).
166. S. Chatrchyan *et al.* (CMS Collab.), Phys. Lett. **B727**, 77 (2013).
167. R. Aaij *et al.* (LHCb Collab.), JHEP **1308**, 131 (2013).
168. V. Khachatryan *et al.* (CMS Collab.), Phys. Lett. **B753**, 424 (2016).
169. R. Aaij *et al.* (LHCb Collab.), JHEP **1602**, 104 (2016).
170. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **111**, 151801 (2013).
171. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **110**, 031801 (2013).
172. R. Aaij *et al.* (LHCb Collab.), JHEP **1207**, 133 (2012).
173. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **111**, 191801 (2013).
174. S. Descotes-Genon, J. Matias, M. Ramon, and J. Virto, JHEP **1301**, 048 (2013).
175. R. Aaij *et al.* (LHCb Collab.), JHEP **1405**, 082 (2014).
176. R. Aaij *et al.* (LHCb Collab.), JHEP **1504**, 064 (2015).
177. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **113**, 151601 (2014).
178. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **107**, 239903 (2011).
179. V. M. Abazov *et al.* (D0 Collab.), Phys. Rev. **D87**, 072006 (2013).
180. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **110**, 021801 (2013).
181. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **111**, 101805 (2013).
182. S. Chatrchyan *et al.* (CMS Collab.), Phys. Rev. Lett. **111**, 101804 (2013).
183. V. Khachatryan *et al.* (CMS and LHCb Collab.), Nature **522**, 68 (2015).
184. M. Aaboud *et al.* (ATLAS Collab.), arXiv:1604.04263 [hep-ex], submitted to Eur. Phys. J. C..
185. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **102**, 201801 (2009).
186. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **111**, 141801 (2013).
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- A NOTE ON HFAG ACTIVITIES**
- Revised September 2015 by T. Gershon (University of Warwick) and A.J. Schwartz (University of Cincinnati)
- The Heavy Flavor Averaging Group (HFAG) is an international collaboration of physicists from experiments measuring properties of heavy flavored particles. HFAG calculates for the HEP community world average values of quantities such as lifetimes, branching fractions, form factors, mixing parameters, and *CP*-violating asymmetries. Most parameters concern decays of *B* and *D* mesons and  $\tau$  leptons, and many are related to elements of the Cabibbo-Kobayashi-Maskawa (CKM) quark mixing matrix [1,2].
- HFAG was originally formed in 2002 to continue the activities of the LEP Heavy Flavor Steering group. Since its inception a wide range of results have become available from increasingly larger data sets, and consequently HFAG has expanded to include seven subgroups. These are as follows:
- *b*-hadron lifetimes and oscillations, including parameters of *CP* violation in *b* mixing;
  - decay-time-dependent *CP* violation in *B* decays, and angles of the CKM Unitarity Triangle;
  - semileptonic decays of *b*-hadrons ( $B \rightarrow X\ell\nu$ ,  $\ell = e, \mu, \tau$ ), including determinations of the CKM matrix elements  $|V_{cb}|$  and  $|V_{ub}|$ ;
  - *b*-hadron decays to hadronic final states containing *c*-quarks (open charm and charmonium);
  - (rarer) *b*-hadron decays to final states not containing *c*-quarks, including fully hadronic, semileptonic ( $B \rightarrow X\ell\ell, X\nu\bar{\nu}$ ), leptonic, and radiative decays;

# Meson Particle Listings

## $b$ -flavored hadrons, $B^\pm$

- $c$ -hadron physics including branching fractions,  $CP$ - and  $T$ -violating asymmetries,  $D^0$ - $\bar{D}^0$  mixing, semi-leptonic decays, and properties of excited  $D$  states;
- $\tau$ -lepton physics including lepton universality tests, determination of the CKM matrix element  $|V_{us}|$ , and searches for lepton flavor violation.

Each subgroup has one or two conveners and typically a half-dozen members representing experiments currently or recently making measurements in that area. Most groups contain representatives from the Belle, BaBar, and LHCb experiments, while some groups contain representatives or contacts from the BESIII, CLEO(c), CDF and DØ experiments. Members of HFAG are appointed by their respective experimental collaborations. There are two co-leaders of HFAG; these were originally appointed by the managements of the BaBar and Belle collaborations and are now appointed by the managements of Belle/Belle II and LHCb.

The averaging procedures used by HFAG are similar to those of the PDG [3], but there are some differences. When calculating world averages, common input parameters used in the different analyses are adjusted (rescaled) to common values. Close communication between representatives of the experiments and HFAG members performing averaging calculations help ensure that measurement uncertainties, known correlations, and systematic effects are properly accounted for. The confidence level of the fit is provided to indicate the consistency of the measurements included in the average. In the case of obtaining a world average with a small confidence level, *i.e.*, a large  $\chi^2$  per degree of freedom, HFAG does not usually scale the resulting uncertainty as the PDG does. Rather, the systematic uncertainties of each input measurement are reviewed with experts from the experiments to better understand the discrepancy. Unless inconsistencies between the measurements are found, no correction is made to the calculated uncertainty. If special treatment is necessary to calculate an average, or in case an approximation used in an average calculation might not be sufficiently accurate (*e.g.*, assuming Gaussian errors when the likelihood function indicates non-Gaussian behavior), a note is included to describe this treatment.

In general, HFAG uses all publicly available results that have written documentation such as a journal publication, preprint or conference note. These include preliminary results presented at conferences or workshops. However, preliminary results that remain unpublished for an extended period of time, or for which no publication is planned, are not included. A special subset of HFAG world averages are included in the PDG Listings. For these averages, the standard fitting procedures are performed but only input measurements that are published or accepted for publication are used. The averages provided by HFAG are listed by the PDG as “OUR EVALUATION” with a corresponding note.

All HFAG world averages and listings of all input measurements are documented in an approximately biennial preprint

posted to the arXiv preprint server; the most recent version is Ref. 4. The latest results and plots are posted on an extensive set of webpages that are updated several times per year; these are available at

<http://www.slac.stanford.edu/xorg/hfag>.

### References:

1. N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
2. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
3. See Section 5 of the “Introduction” to this *Review*.
4. Y. Amhis *et al.* [HFAG Collab.], [arXiv:1412.7515](https://arxiv.org/abs/1412.7515).

$B^\pm$

$$I(J^P) = \frac{1}{2}(0^-)$$

Quantum numbers not measured. Values shown are quark-model predictions.

See also the  $B^\pm/B^0$  ADMIXTURE and  $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE sections.

### $B^\pm$ MASS

The fit uses  $m_{B^\pm}$ ,  $(m_{B^0} - m_{B^\pm})$ , and  $m_{B^0}$  to determine  $m_{B^\pm}$ ,  $m_{B^0}$ , and the mass difference.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5279.31 ± 0.15 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>5279.25 ± 0.26 OUR AVERAGE</b>				
5279.38 ± 0.11 ± 0.33		<sup>1</sup> AAIJ	12E	LHCb $pp$ at 7 TeV
5279.10 ± 0.41 ± 0.36		<sup>2</sup> ACOSTA	06	CDF $p\bar{p}$ at 1.96 TeV
5279.1 ± 0.4 ± 0.4	526	<sup>3</sup> CSORNA	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
5279.1 ± 1.7 ± 1.4	147	ABE	96B	CDF $p\bar{p}$ at 1.8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5278.8 ± 0.54 ± 2.0	362	ALAM	94	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
5278.3 ± 0.4 ± 2.0		BORTOLETTO	092	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
5280.5 ± 1.0 ± 2.0		<sup>4</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
5275.8 ± 1.3 ± 3.0	32	ALBRECHT	87C	ARG $e^+e^- \rightarrow \Upsilon(4S)$
5278.2 ± 1.8 ± 3.0	12	<sup>5</sup> ALBRECHT	87D	ARG $e^+e^- \rightarrow \Upsilon(4S)$
5278.6 ± 0.8 ± 2.0		BEBEK	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B^+ \rightarrow J/\psi K^+$  fully reconstructed decays.

<sup>2</sup> Uses exclusively reconstructed final states containing a  $J/\psi \rightarrow \mu^+\mu^-$  decays.

<sup>3</sup> CSORNA 00 uses fully reconstructed 526  $B^+ \rightarrow J/\psi(\Upsilon) K^+$  events and invariant masses without beam constraint.

<sup>4</sup> ALBRECHT 90J assumes 10580 for  $\Upsilon(4S)$  mass. Supersedes ALBRECHT 87C and ALBRECHT 87D.

<sup>5</sup> Found using fully reconstructed decays with  $J/\psi(1S)$ . ALBRECHT 87D assume  $m_{\Upsilon(4S)} = 10577$  MeV.

### $B^\pm$ MEAN LIFE

See  $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE section for data on  $B$ -hadron mean life averaged over species of bottom particles.

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

VALUE ( $10^{-12}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.638 ± 0.004 OUR EVALUATION</b>				
1.637 ± 0.004 ± 0.003		AAIJ	14E	LHCb $pp$ at 7 TeV
1.639 ± 0.009 ± 0.009		<sup>1</sup> AALTONEN	11	CDF $p\bar{p}$ at 1.96 TeV
1.663 ± 0.023 ± 0.015		<sup>2</sup> AALTONEN	11B	CDF $p\bar{p}$ at 1.96 TeV
1.635 ± 0.011 ± 0.011		<sup>3</sup> ABE	05B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
1.624 ± 0.014 ± 0.018		<sup>4</sup> ABDALLAH	04E	DLPH $e^+e^- \rightarrow Z$
1.636 ± 0.058 ± 0.025		<sup>5</sup> ACOSTA	02C	CDF $p\bar{p}$ at 1.8 TeV
1.673 ± 0.032 ± 0.023		<sup>6</sup> AUBERT	01F	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.648 ± 0.049 ± 0.035		<sup>7</sup> BARATE	00R	ALEP $e^+e^- \rightarrow Z$

1.643 ± 0.037 ± 0.025	8	ABBIENDI	99J	OPAL	$e^+e^- \rightarrow Z$
1.637 ± 0.058 ± 0.045 0.043	7	ABE	98Q	CDF	$p\bar{p}$ at 1.8 TeV
1.66 ± 0.06 ± 0.03	8	ACCIARRI	98S	L3	$e^+e^- \rightarrow Z$
1.66 ± 0.06 ± 0.05	8	ABE	97J	SLD	$e^+e^- \rightarrow Z$
1.58 +0.21 +0.04 -0.18 -0.03	94	5	BUSKULIC	96J	ALEP $e^+e^- \rightarrow Z$
1.61 ± 0.16 ± 0.12	7.9	ABREU	95Q	DLPH	$e^+e^- \rightarrow Z$
1.72 ± 0.08 ± 0.06	10	ADAM	95	DLPH	$e^+e^- \rightarrow Z$
1.52 ± 0.14 ± 0.09	7	AKERS	95T	OPAL	$e^+e^- \rightarrow Z$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.695 ± 0.026 ± 0.015	6	ABE	02H	BELL	Repl. by ABE 05B
1.68 ± 0.07 ± 0.02	5	ABE	98B	CDF	Repl. by ACOSTA 02C
1.56 ± 0.13 ± 0.06	7	ABE	96C	CDF	Repl. by ABE 98Q
1.58 ± 0.09 ± 0.03	11	BUSKULIC	96J	ALEP	$e^+e^- \rightarrow Z$
1.58 ± 0.09 ± 0.04	7	BUSKULIC	96J	ALEP	Repl. by BARATE 00R
1.70 ± 0.09	12	ADAM	95	DLPH	$e^+e^- \rightarrow Z$
1.61 ± 0.16 ± 0.05	148	5	ABE	94D	CDF Repl. by ABE 98B
1.30 +0.33 +0.29 -0.29 ± 0.16	92	7	ABREU	93D	DLPH Sup. by ABREU 95Q
1.56 ± 0.19 ± 0.13	134	10	ABREU	93G	DLPH Sup. by ADAM 95
1.51 +0.30 +0.12 -0.28 -0.14	59	7	ACTON	93C	OPAL Sup. by AKERS 95T
1.47 +0.22 +0.15 -0.19 -0.14	77	7	BUSKULIC	93D	ALEP Sup. by BUSKULIC 96J

<sup>1</sup> Measured mean life using fully reconstructed decays ( $J/\psi K^{(*)}$ ).

<sup>2</sup> Measured using  $B^- \rightarrow D^0 \pi^-$  with  $D^0 \rightarrow K^- \pi^+$  events that were selected using a silicon vertex trigger.

<sup>3</sup> Measurement performed using a combined fit of  $CP$ -violation, mixing and lifetimes.

<sup>4</sup> Measurement performed using an inclusive reconstruction and  $B$  flavor identification technique.

<sup>5</sup> Measured mean life using fully reconstructed decays.

<sup>6</sup> Events are selected in which one  $B$  meson is fully reconstructed while the second  $B$  meson is reconstructed inclusively.

<sup>7</sup> Data analyzed using  $D/D^* \ell X$  event vertices.

<sup>8</sup> Data analyzed using charge of secondary vertex.

<sup>9</sup> ABREU 95Q assumes  $B(B^0 \rightarrow D^{*-} \ell^+ \nu_\ell) = 3.2 \pm 1.7\%$ .

<sup>10</sup> Data analyzed using vertex-charge technique to tag  $B$  charge.

<sup>11</sup> Combined result of  $D/D^* \ell X$  analysis and fully reconstructed  $B$  analysis.

<sup>12</sup> Combined ABREU 95Q and ADAM 95 result.

### $\tau_{B^+}/\tau_{B^-}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.002 ± 0.004 ± 0.002</b>	1	AAIJ	14E LHCB $p\bar{p}$ at 7 TeV

<sup>1</sup> Measured using  $B^\pm \rightarrow J/\psi K^\pm$  decays.

### $B^+$ DECAY MODES

$B^-$  modes are charge conjugates of the modes below. Modes which do not identify the charge state of the  $B$  are listed in the  $B^\pm/B^0$  ADMIXTURE section.

The branching fractions listed below assume 50%  $B^0 \bar{B}^0$  and 50%  $B^+ B^-$  production at the  $\Upsilon(4S)$ . We have attempted to bring older measurements up to date by rescaling their assumed  $\Upsilon(4S)$  production ratio to 50:50 and their assumed  $D, D_s, D^*$ , and  $\psi$  branching ratios to current values whenever this would affect our averages and best limits significantly.

Indentation is used to indicate a subchannel of a previous reaction. All resonant subchannels have been corrected for resonance branching fractions to the final state so the sum of the subchannel branching fractions can exceed that of the final state.

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm$  anything, the values usually are multiplicities, not branching fractions. They can be greater than one.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Semileptonic and leptonic modes</b>		
$\Gamma_1$ $\ell^+ \nu_\ell$ anything	[a] ( 10.99 ± 0.28 ) %	
$\Gamma_2$ $e^+ \nu_e X_c$	( 10.8 ± 0.4 ) %	
$\Gamma_3$ $D \ell^+ \nu_\ell$ anything	( 9.8 ± 0.7 ) %	
$\Gamma_4$ $\bar{D}^0 \ell^+ \nu_\ell$	[a] ( 2.27 ± 0.11 ) %	
$\Gamma_5$ $\bar{D}^0 \tau^+ \nu_\tau$	( 7.7 ± 2.5 ) × 10 <sup>-3</sup>	
$\Gamma_6$ $\bar{D}^*(2007)^0 \ell^+ \nu_\ell$	[a] ( 5.69 ± 0.19 ) %	
$\Gamma_7$ $\bar{D}^*(2007)^0 \tau^+ \nu_\tau$	( 1.88 ± 0.20 ) %	
$\Gamma_8$ $D^- \pi^+ \ell^+ \nu_\ell$	( 4.2 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_9$ $\bar{D}_0^*(2420)^0 \ell^+ \nu_\ell, \bar{D}_0^{*0} \rightarrow$	( 2.5 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{10}$ $\bar{D}_2^{*-} \pi^+ \ell^+ \nu_\ell, \bar{D}_2^{*0} \rightarrow$	( 1.53 ± 0.16 ) × 10 <sup>-3</sup>	
$\Gamma_{11}$ $D^{(*)} n \pi \ell^+ \nu_\ell (n \geq 1)$	( 1.87 ± 0.26 ) %	
$\Gamma_{12}$ $D^{*-} \pi^+ \ell^+ \nu_\ell$	( 6.1 ± 0.6 ) × 10 <sup>-3</sup>	
$\Gamma_{13}$ $\bar{D}_1^*(2420)^0 \ell^+ \nu_\ell, \bar{D}_1^0 \rightarrow$	( 3.03 ± 0.20 ) × 10 <sup>-3</sup>	
$D^{*-} \pi^+$		

$\Gamma_{14}$ $\bar{D}_1^*(2430)^0 \ell^+ \nu_\ell, \bar{D}_1^{*0} \rightarrow$	( 2.7 ± 0.6 ) × 10 <sup>-3</sup>	
$\Gamma_{15}$ $\bar{D}_2^{*-} \pi^+ \ell^+ \nu_\ell, \bar{D}_2^{*0} \rightarrow$	( 1.01 ± 0.24 ) × 10 <sup>-3</sup>	S=2.0
$\Gamma_{16}$ $\bar{D}^0 \pi^+ \pi^- \ell^+ \nu_\ell$	( 1.6 ± 0.4 ) × 10 <sup>-3</sup>	
$\Gamma_{17}$ $\bar{D}^{*0} \pi^+ \pi^- \ell^+ \nu_\ell$	( 8 ± 5 ) × 10 <sup>-4</sup>	
$\Gamma_{18}$ $D_s^{*-} K^+ \ell^+ \nu_\ell$	( 6.1 ± 1.0 ) × 10 <sup>-4</sup>	
$\Gamma_{19}$ $D_s^- K^+ \ell^+ \nu_\ell$	( 3.0 ± 1.4 ) × 10 <sup>-4</sup>	
$\Gamma_{20}$ $D_s^{*-} K^+ \ell^+ \nu_\ell$	( 2.9 ± 1.9 ) × 10 <sup>-4</sup>	
$\Gamma_{21}$ $\pi^0 \ell^+ \nu_\ell$	( 7.80 ± 0.27 ) × 10 <sup>-5</sup>	
$\Gamma_{22}$ $\pi^0 e^+ \nu_e$		
$\Gamma_{23}$ $\eta \ell^+ \nu_\ell$	( 3.8 ± 0.6 ) × 10 <sup>-5</sup>	
$\Gamma_{24}$ $\eta' \ell^+ \nu_\ell$	( 2.3 ± 0.8 ) × 10 <sup>-5</sup>	
$\Gamma_{25}$ $\omega \ell^+ \nu_\ell$	[a] ( 1.19 ± 0.09 ) × 10 <sup>-4</sup>	
$\Gamma_{26}$ $\omega \mu^+ \nu_\mu$		
$\Gamma_{27}$ $\rho^0 \ell^+ \nu_\ell$	[a] ( 1.58 ± 0.11 ) × 10 <sup>-4</sup>	
$\Gamma_{28}$ $p\bar{p} \ell^+ \nu_\ell$	( 5.8 ± 2.6 ) × 10 <sup>-6</sup>	
$\Gamma_{29}$ $p\bar{p} \mu^+ \nu_\mu$	< 8.5 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{30}$ $p\bar{p} e^+ \nu_e$	( 8.2 ± 4.0 ) × 10 <sup>-6</sup>	
$\Gamma_{31}$ $e^+ \nu_e$	< 9.8 × 10 <sup>-7</sup>	CL=90%
$\Gamma_{32}$ $\mu^+ \nu_\mu$	< 1.0 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{33}$ $\tau^+ \nu_\tau$	( 1.09 ± 0.24 ) × 10 <sup>-4</sup>	S=1.2
$\Gamma_{34}$ $\ell^+ \nu_\ell \gamma$	< 3.5 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{35}$ $e^+ \nu_e \gamma$	< 6.1 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{36}$ $\mu^+ \nu_\mu \gamma$	< 3.4 × 10 <sup>-6</sup>	CL=90%

### Inclusive modes

$\Gamma_{37}$ $D^0 X$	( 8.6 ± 0.7 ) %
$\Gamma_{38}$ $\bar{D}^0 X$	( 79 ± 4 ) %
$\Gamma_{39}$ $D^+ X$	( 2.5 ± 0.5 ) %
$\Gamma_{40}$ $D^- X$	( 9.9 ± 1.2 ) %
$\Gamma_{41}$ $D_s^+ X$	( 7.9 ± 1.3 ) %
$\Gamma_{42}$ $D_s^- X$	( 1.10 ± 0.40 ) % 0.32
$\Gamma_{43}$ $\Lambda_c^+ X$	( 2.1 ± 0.9 ) % 0.6
$\Gamma_{44}$ $\bar{\Lambda}_c^- X$	( 2.8 ± 1.1 ) % 0.9
$\Gamma_{45}$ $\bar{c} X$	( 97 ± 4 ) %
$\Gamma_{46}$ $c X$	( 23.4 ± 2.2 ) % 1.8
$\Gamma_{47}$ $c/\bar{c} X$	( 120 ± 6 ) %

### $D, D^*,$ or $D_s$ modes

$\Gamma_{48}$ $\bar{D}^0 \pi^+$	( 4.80 ± 0.15 ) × 10 <sup>-3</sup>
$\Gamma_{49}$ $D_{CP(+1)} \pi^+$	[b] ( 2.19 ± 0.24 ) × 10 <sup>-3</sup>
$\Gamma_{50}$ $D_{CP(-1)} \pi^+$	[b] ( 2.1 ± 0.4 ) × 10 <sup>-3</sup>
$\Gamma_{51}$ $\bar{D}^0 \rho^+$	( 1.34 ± 0.18 ) %
$\Gamma_{52}$ $\bar{D}^0 K^+$	( 3.69 ± 0.17 ) × 10 <sup>-4</sup>
$\Gamma_{53}$ $D_{CP(+1)} K^+$	[b] ( 1.91 ± 0.14 ) × 10 <sup>-4</sup>
$\Gamma_{54}$ $D_{CP(-1)} K^+$	[b] ( 1.99 ± 0.19 ) × 10 <sup>-4</sup>
$\Gamma_{55}$ $[K^- \pi^+]_D K^+$	[c] < 2.8 × 10 <sup>-7</sup>
$\Gamma_{56}$ $[K^+ \pi^-]_D K^+$	[c] < 1.8 × 10 <sup>-5</sup>
$\Gamma_{57}$ $[K^- \pi^+ \pi^0]_D K^+$	seen
$\Gamma_{58}$ $[K^+ \pi^- \pi^0]_D K^+$	seen
$\Gamma_{59}$ $[K^- \pi^+ \pi^+ \pi^-]_D K^+$	seen
$\Gamma_{60}$ $[K^+ \pi^- \pi^+ \pi^-]_D K^+$	seen
$\Gamma_{61}$ $[K^- \pi^+]_D K^*(892)^+$	[c]
$\Gamma_{62}$ $[K^+ \pi^-]_D K^*(892)^+$	[c]
$\Gamma_{63}$ $[K^- \pi^+]_D \pi^+$	[c] ( 6.3 ± 1.1 ) × 10 <sup>-7</sup>
$\Gamma_{64}$ $[K^+ \pi^-]_D \pi^+$	( 1.68 ± 0.31 ) × 10 <sup>-4</sup>
$\Gamma_{65}$ $[K^- \pi^+ \pi^0]_D \pi^+$	seen
$\Gamma_{66}$ $[K^+ \pi^- \pi^0]_D \pi^+$	seen
$\Gamma_{67}$ $[K^- \pi^+ \pi^+ \pi^-]_D \pi^+$	seen
$\Gamma_{68}$ $[K^+ \pi^- \pi^+ \pi^-]_D \pi^+$	seen
$\Gamma_{69}$ $[K^- \pi^+]_{(D\pi)} \pi^+$	
$\Gamma_{70}$ $[K^+ \pi^-]_{(D\pi)} \pi^+$	
$\Gamma_{71}$ $[K^- \pi^+]_{(D\gamma)} \pi^+$	
$\Gamma_{72}$ $[K^+ \pi^-]_{(D\gamma)} \pi^+$	
$\Gamma_{73}$ $[K^- \pi^+]_{(D\pi)} K^+$	
$\Gamma_{74}$ $[K^+ \pi^-]_{(D\pi)} K^+$	
$\Gamma_{75}$ $[K^- \pi^+]_{(D\gamma)} K^+$	
$\Gamma_{76}$ $[K^+ \pi^-]_{(D\gamma)} K^+$	

## Meson Particle Listings

 $B^\pm$ 

$\Gamma_{77}$	$[\pi^+\pi^-\pi^0]_D K^-$	( 4.6 $\pm$ 0.9 ) $\times 10^{-6}$		$\Gamma_{140}$	$\overline{D}_2^*(2462)^0 \pi^+ \times B(\overline{D}_2^{*0} \rightarrow D^{*0}(2010)^- \pi^+)$	( 2.2 $\pm$ 1.1 ) $\times 10^{-4}$
$\Gamma_{78}$	$[K_S^0 K^+ \pi^-]_D K^+$	seen		$\Gamma_{141}$	$\overline{D}_0^*(2400)^0 \pi^+ \times B(\overline{D}_0^*(2400)^0 \rightarrow D^- \pi^+)$	( 6.4 $\pm$ 1.4 ) $\times 10^{-4}$
$\Gamma_{79}$	$[K_S^0 K^- \pi^+]_D K^+$	seen		$\Gamma_{142}$	$\overline{D}_1(2421)^0 \pi^+ \times B(\overline{D}_1(2421)^0 \rightarrow D^{*-} \pi^+)$	( 6.8 $\pm$ 1.5 ) $\times 10^{-4}$
$\Gamma_{80}$	$[K^*(892)^+ K^-]_D K^+$	seen		$\Gamma_{143}$	$\overline{D}_2^*(2462)^0 \pi^+ \times B(\overline{D}_2^*(2462)^0 \rightarrow D^{*-} \pi^+)$	( 1.8 $\pm$ 0.5 ) $\times 10^{-4}$
$\Gamma_{81}$	$[K_S^0 K^- \pi^+]_D \pi^+$	seen		$\Gamma_{144}$	$\overline{D}_1(2427)^0 \pi^+ \times B(\overline{D}_1(2427)^0 \rightarrow D^{*-} \pi^+)$	( 5.0 $\pm$ 1.2 ) $\times 10^{-4}$
$\Gamma_{82}$	$[K^*(892)^+ K^-]_D \pi^+$	seen		$\Gamma_{145}$	$\overline{D}_1(2420)^0 \pi^+ \times B(\overline{D}_1^0 \rightarrow \overline{D}^{*0} \pi^+ \pi^-)$	< 6 $\times 10^{-6}$ CL=90%
$\Gamma_{83}$	$[K_S^0 K^+ \pi^-]_D \pi^+$	seen		$\Gamma_{146}$	$\overline{D}_1^*(2420)^0 \rho^+$	< 1.4 $\times 10^{-3}$ CL=90%
$\Gamma_{84}$	$[K^*(892)^- K^+]_D \pi^+$	seen		$\Gamma_{147}$	$\overline{D}_2^*(2460)^0 \pi^+$	< 1.3 $\times 10^{-3}$ CL=90%
$\Gamma_{85}$	$[K^+ K^- \pi^0]_D K^+$			$\Gamma_{148}$	$\overline{D}_2^*(2460)^0 \pi^+ \times B(\overline{D}_2^{*0} \rightarrow \overline{D}^{*0} \pi^+ \pi^-)$	< 2.2 $\times 10^{-5}$ CL=90%
$\Gamma_{86}$	$[K^+ K^- \pi^0]_D \pi^+$			$\Gamma_{149}$	$\overline{D}_2^*(2460)^0 \rho^+$	< 4.7 $\times 10^{-3}$ CL=90%
$\Gamma_{87}$	$[\pi^+ \pi^- \pi^0]_D K^+$			$\Gamma_{150}$	$\overline{D}^0 D_s^+$	( 9.0 $\pm$ 0.9 ) $\times 10^{-3}$
$\Gamma_{88}$	$[\pi^+ \pi^- \pi^0]_D \pi^+$			$\Gamma_{151}$	$D_{s0}^*(2317)^+ \overline{D}^0, D_{s0}^{*+} \rightarrow D_s^+ \pi^0$	( 7.9 $\pm$ 1.5 $\pm$ 1.3 ) $\times 10^{-4}$
$\Gamma_{89}$	$\overline{D}^0 K^*(892)^+$	( 5.3 $\pm$ 0.4 ) $\times 10^{-4}$		$\Gamma_{152}$	$D_{s0}(2317)^+ \overline{D}^0 \times B(D_{s0}(2317)^+ \rightarrow D_s^{*+} \gamma)$	< 7.6 $\times 10^{-4}$ CL=90%
$\Gamma_{90}$	$D_{CP(-1)} K^*(892)^+$	[b] ( 2.7 $\pm$ 0.8 ) $\times 10^{-4}$		$\Gamma_{153}$	$D_{s0}(2317)^+ \overline{D}^*(2007)^0 \times B(D_{s0}(2317)^+ \rightarrow D_s^+ \pi^0)$	( 9 $\pm$ 7 ) $\times 10^{-4}$
$\Gamma_{91}$	$D_{CP(+1)} K^*(892)^+$	[b] ( 5.8 $\pm$ 1.1 ) $\times 10^{-4}$		$\Gamma_{154}$	$D_{sJ}(2457)^+ \overline{D}^0$	( 3.1 $\pm$ 1.0 $\pm$ 0.9 ) $\times 10^{-3}$
$\Gamma_{92}$	$\overline{D}^0 K^+ \pi^+ \pi^-$	( 5.4 $\pm$ 2.2 ) $\times 10^{-4}$		$\Gamma_{155}$	$D_{sJ}(2457)^+ \overline{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \gamma)$	( 4.6 $\pm$ 1.3 $\pm$ 1.1 ) $\times 10^{-4}$
$\Gamma_{93}$	$[K^+ \pi^-]_D K^+ \pi^- \pi^+$			$\Gamma_{156}$	$D_{sJ}(2457)^+ \overline{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \pi^0)$	< 2.2 $\times 10^{-4}$ CL=90%
$\Gamma_{94}$	$[K^- \pi^+]_D K^+ \pi^- \pi^+$			$\Gamma_{157}$	$D_{sJ}(2457)^+ \overline{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \pi^0)$	< 2.7 $\times 10^{-4}$ CL=90%
$\Gamma_{95}$	$D_{CP(+1)} K^+ \pi^- \pi^+$			$\Gamma_{158}$	$D_{sJ}(2457)^+ \overline{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^{*+} \gamma)$	< 9.8 $\times 10^{-4}$ CL=90%
$\Gamma_{96}$	$\overline{D}^0 K^+ \overline{K}^0$	( 5.5 $\pm$ 1.6 ) $\times 10^{-4}$		$\Gamma_{159}$	$D_{sJ}(2457)^+ \overline{D}^*(2007)^0$	( 1.20 $\pm$ 0.30 ) %
$\Gamma_{97}$	$\overline{D}^0 K^+ \overline{K}^*(892)^0$	( 7.5 $\pm$ 1.7 ) $\times 10^{-4}$		$\Gamma_{160}$	$D_{sJ}(2457)^+ \overline{D}^*(2007)^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \gamma)$	( 1.4 $\pm$ 0.7 $\pm$ 0.6 ) $\times 10^{-3}$
$\Gamma_{98}$	$\overline{D}^0 \pi^+ \pi^+ \pi^-$	( 5.7 $\pm$ 2.2 ) $\times 10^{-3}$	S=3.6	$\Gamma_{161}$	$\overline{D}^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+ + D^*(2010)^+ K^0)$	( 4.0 $\pm$ 1.0 ) $\times 10^{-4}$
$\Gamma_{99}$	$[K^- \pi^+]_D \pi^+ \pi^- \pi^+$			$\Gamma_{162}$	$\overline{D}^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+ + D^*(2010)^+ K^0)$	( 2.2 $\pm$ 0.7 ) $\times 10^{-4}$
$\Gamma_{100}$	$\overline{D}^0 \pi^+ \pi^+ \pi^-$ nonresonant	( 5 $\pm$ 4 ) $\times 10^{-3}$		$\Gamma_{163}$	$\overline{D}^*(2007)^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+ + D^*(2010)^+ K^0)$	( 5.5 $\pm$ 1.6 ) $\times 10^{-4}$
$\Gamma_{101}$	$\overline{D}^0 \pi^+ \rho^0$	( 4.2 $\pm$ 3.0 ) $\times 10^{-3}$		$\Gamma_{164}$	$\overline{D}^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^{*+} K^0)$	( 2.3 $\pm$ 1.1 ) $\times 10^{-4}$
$\Gamma_{102}$	$\overline{D}^0 a_1(1260)^+$	( 4 $\pm$ 4 ) $\times 10^{-3}$		$\Gamma_{165}$	$\overline{D}^0 D_{sJ}(2700)^+ \times B(D_{sJ}(2700)^+ \rightarrow D^0 K^+)$	( 5.6 $\pm$ 1.8 ) $\times 10^{-4}$ S=1.7
$\Gamma_{103}$	$\overline{D}^0 \omega \pi^+$	( 4.1 $\pm$ 0.9 ) $\times 10^{-3}$		$\Gamma_{166}$	$\overline{D}^{*0} D_{s1}(2536)^+ \rightarrow D_s^{*+} \pi^0$	( 3.9 $\pm$ 2.6 ) $\times 10^{-4}$
$\Gamma_{104}$	$D^*(2010)^- \pi^+ \pi^+$	( 1.35 $\pm$ 0.22 ) $\times 10^{-3}$		$\Gamma_{167}$	$\overline{D}^0 D_{sJ}(2573)^+, D_{sJ}^+ \rightarrow D^0 K^+$	( 8 $\pm$ 15 ) $\times 10^{-6}$
$\Gamma_{105}$	$\overline{D}_1(2420)^0 \pi^+, \overline{D}_1^- \rightarrow D^*(2010)^- \pi^+$	( 5.3 $\pm$ 2.3 ) $\times 10^{-4}$		$\Gamma_{168}$	$\overline{D}^{*0} D_{sJ}(2573), D_{sJ}^+ \rightarrow D^0 K^+$	< 2 $\times 10^{-4}$ CL=90%
$\Gamma_{106}$	$D^- \pi^+ \pi^+$	( 1.07 $\pm$ 0.05 ) $\times 10^{-3}$		$\Gamma_{169}$	$\overline{D}^*(2007)^0 D_{sJ}(2573), D_{sJ}^+ \rightarrow D^0 K^+$	< 5 $\times 10^{-4}$ CL=90%
$\Gamma_{107}$	$D^- K^+ \pi^+$	( 7.7 $\pm$ 0.5 ) $\times 10^{-5}$		$\Gamma_{170}$	$\overline{D}^0 D_s^{*+}$	( 7.6 $\pm$ 1.6 ) $\times 10^{-3}$
$\Gamma_{108}$	$D_0^*(2400)^0 K^+, D_0^{*0} \rightarrow D^- \pi^+$	( 6.1 $\pm$ 2.4 ) $\times 10^{-4}$		$\Gamma_{171}$	$\overline{D}^*(2007)^0 D_s^+$	( 8.2 $\pm$ 1.7 ) $\times 10^{-3}$
$\Gamma_{109}$	$D_1^*(2760)^0 K^+, D_1^{*0} \rightarrow D^- \pi^+$	( 3.6 $\pm$ 1.2 ) $\times 10^{-4}$		$\Gamma_{172}$	$\overline{D}^*(2007)^0 D_s^{*+}$	( 1.71 $\pm$ 0.24 ) %
$\Gamma_{110}$	$D_2^*(2460)^0 K^+, D_2^{*0} \rightarrow \rho^- \pi^+$	( 2.32 $\pm$ 0.23 ) $\times 10^{-3}$		$\Gamma_{173}$	$D_s^{(*)+} \overline{D}^{*0}$	( 2.7 $\pm$ 1.2 ) %
$\Gamma_{111}$	$D^+ K^0$	< 2.9 $\times 10^{-6}$ CL=90%		$\Gamma_{174}$	$\overline{D}^*(2007)^0 D^*(2010)^+$	( 8.1 $\pm$ 1.7 ) $\times 10^{-4}$
$\Gamma_{112}$	$D^+ K^{*0}$	< 1.8 $\times 10^{-6}$ CL=90%		$\Gamma_{175}$	$\overline{D}^0 D^*(2010)^+ + \overline{D}^*(2007)^0 D^+$	< 1.30 % CL=90%
$\Gamma_{113}$	$D^+ \overline{K}^{*0}$	< 1.4 $\times 10^{-6}$ CL=90%		$\Gamma_{176}$	$\overline{D}^0 D^*(2010)^+$	( 3.9 $\pm$ 0.5 ) $\times 10^{-4}$
$\Gamma_{114}$	$\overline{D}^*(2007)^0 \pi^+$	( 5.18 $\pm$ 0.26 ) $\times 10^{-3}$		$\Gamma_{177}$	$\overline{D}^0 D^+$	( 3.8 $\pm$ 0.4 ) $\times 10^{-4}$
$\Gamma_{115}$	$\overline{D}_{CP(+1)}^{*0} \pi^+$	[d] ( 2.9 $\pm$ 0.7 ) $\times 10^{-3}$		$\Gamma_{178}$	$\overline{D}^0 D^+ K^0$	( 1.55 $\pm$ 0.21 ) $\times 10^{-3}$
$\Gamma_{116}$	$\overline{D}_{CP(-1)}^{*0} \pi^+$	[d] ( 2.6 $\pm$ 1.0 ) $\times 10^{-3}$		$\Gamma_{179}$	$D^+ \overline{D}^*(2007)^0$	( 6.3 $\pm$ 1.7 ) $\times 10^{-4}$
$\Gamma_{117}$	$\overline{D}^*(2007)^0 \omega \pi^+$	( 4.5 $\pm$ 1.2 ) $\times 10^{-3}$		$\Gamma_{180}$	$\overline{D}^*(2007)^0 D^+ K^0$	( 2.1 $\pm$ 0.5 ) $\times 10^{-3}$
$\Gamma_{118}$	$\overline{D}^*(2007)^0 \rho^+$	( 9.8 $\pm$ 1.7 ) $\times 10^{-3}$				
$\Gamma_{119}$	$\overline{D}^*(2007)^0 K^+$	( 4.20 $\pm$ 0.34 ) $\times 10^{-4}$				
$\Gamma_{120}$	$\overline{D}_{CP(+1)}^{*0} K^+$	[d] ( 2.8 $\pm$ 0.4 ) $\times 10^{-4}$				
$\Gamma_{121}$	$\overline{D}_{CP(-1)}^{*0} K^+$	[d] ( 2.31 $\pm$ 0.33 ) $\times 10^{-4}$				
$\Gamma_{122}$	$\overline{D}^*(2007)^0 K^*(892)^+$	( 8.1 $\pm$ 1.4 ) $\times 10^{-4}$				
$\Gamma_{123}$	$\overline{D}^*(2007)^0 K^+ \overline{K}^0$	< 1.06 $\times 10^{-3}$ CL=90%				
$\Gamma_{124}$	$\overline{D}^*(2007)^0 K^+ K^*(892)^0$	( 1.5 $\pm$ 0.4 ) $\times 10^{-3}$				
$\Gamma_{125}$	$\overline{D}^*(2007)^0 \pi^+ \pi^+ \pi^-$	( 1.03 $\pm$ 0.12 ) %				
$\Gamma_{126}$	$\overline{D}^*(2007)^0 a_1(1260)^+$	( 1.9 $\pm$ 0.5 ) %				
$\Gamma_{127}$	$\overline{D}^*(2007)^0 \pi^- \pi^+ \pi^+ \pi^0$	( 1.8 $\pm$ 0.4 ) %				
$\Gamma_{128}$	$\overline{D}^{*0} 3\pi^+ 2\pi^-$	( 5.7 $\pm$ 1.2 ) $\times 10^{-3}$				
$\Gamma_{129}$	$D^*(2010)^+ \pi^0$	< 3.6 $\times 10^{-6}$				
$\Gamma_{130}$	$D^*(2010)^+ K^0$	< 9.0 $\times 10^{-6}$ CL=90%				
$\Gamma_{131}$	$D^*(2010)^- \pi^+ \pi^+ \pi^0$	( 1.5 $\pm$ 0.7 ) %				
$\Gamma_{132}$	$D^*(2010)^- \pi^+ \pi^+ \pi^+ \pi^-$	( 2.6 $\pm$ 0.4 ) $\times 10^{-3}$				
$\Gamma_{133}$	$\overline{D}^{*0} \pi^+$	[e] ( 5.9 $\pm$ 1.3 ) $\times 10^{-3}$				
$\Gamma_{134}$	$\overline{D}_1^*(2420)^0 \pi^+$	( 1.5 $\pm$ 0.6 ) $\times 10^{-3}$	S=1.3			
$\Gamma_{135}$	$\overline{D}_1(2420)^0 \pi^+ \times B(\overline{D}_1^0 \rightarrow \overline{D}^0 \pi^+ \pi^-)$	( 2.5 $\pm$ 1.6 $\pm$ 1.4 ) $\times 10^{-4}$	S=4.0			
$\Gamma_{136}$	$\overline{D}_1(2420)^0 \pi^+ \times B(\overline{D}_1^0 \rightarrow \overline{D}^0 \pi^+ \pi^- \text{ (nonresonant)})$	( 2.3 $\pm$ 1.0 ) $\times 10^{-4}$				
$\Gamma_{137}$	$\overline{D}_2^*(2462)^0 \pi^+ \times B(\overline{D}_2^*(2462)^0 \rightarrow D^- \pi^+)$	( 3.5 $\pm$ 0.4 ) $\times 10^{-4}$				
$\Gamma_{138}$	$\overline{D}_2^*(2462)^0 \pi^+ \times B(\overline{D}_2^{*0} \rightarrow \overline{D}^0 \pi^- \pi^+)$	( 2.3 $\pm$ 1.1 ) $\times 10^{-4}$				
$\Gamma_{139}$	$\overline{D}_2^*(2462)^0 \pi^+ \times B(\overline{D}_2^{*0} \rightarrow \overline{D}^0 \pi^- \pi^+ \text{ (nonresonant)})$	< 1.7 $\times 10^{-4}$ CL=90%				



$\Gamma_{181} \bar{D}^0 D^*(2010)^+ K^0$	$(3.8 \pm 0.4) \times 10^{-3}$		$\Gamma_{246} X(3872) K^*(892)^+, X \rightarrow J/\psi \gamma$	$< 4.8$	$\times 10^{-6}$ CL=90%
$\Gamma_{182} \bar{D}^*(2007)^0 D^*(2010)^+ K^0$	$(9.2 \pm 1.2) \times 10^{-3}$		$\Gamma_{247} X(3872) K^*(892)^+, X \rightarrow \psi(2S) \gamma$	$< 2.8$	$\times 10^{-5}$ CL=90%
$\Gamma_{183} \bar{D}^0 D^0 K^+$	$(1.45 \pm 0.33) \times 10^{-3}$	S=2.6	$\Gamma_{248} X(3872)^+ K^0, X^+ \rightarrow J/\psi(1S) \pi^+ \pi^0$	$[f] < 6.1$	$\times 10^{-6}$ CL=90%
$\Gamma_{184} \bar{D}^*(2007)^0 D^0 K^+$	$(2.26 \pm 0.23) \times 10^{-3}$		$\Gamma_{249} X(3872) K^0 \pi^+, X \rightarrow J/\psi(1S) \pi^+ \pi^-$	$(1.06 \pm 0.31) \times 10^{-5}$	
$\Gamma_{185} \bar{D}^0 D^*(2007)^0 K^+$	$(6.3 \pm 0.5) \times 10^{-3}$		$\Gamma_{250} X(4430)^+ K^0, X^+ \rightarrow J/\psi \pi^+$	$< 1.5$	$\times 10^{-5}$ CL=95%
$\Gamma_{186} \bar{D}^*(2007)^0 D^*(2007)^0 K^+$	$(1.12 \pm 0.13) \%$		$\Gamma_{251} X(4430)^+ K^0, X^+ \rightarrow \psi(2S) \pi^+$	$< 4.7$	$\times 10^{-5}$ CL=95%
$\Gamma_{187} D^- D^+ K^+$	$(2.2 \pm 0.7) \times 10^{-4}$		$\Gamma_{252} X(4260)^0 K^+, X^0 \rightarrow J/\psi \pi^+ \pi^-$	$< 2.9$	$\times 10^{-5}$ CL=95%
$\Gamma_{188} D^- D^*(2010)^+ K^+$	$(6.3 \pm 1.1) \times 10^{-4}$		$\Gamma_{253} X(3915) K^+, X \rightarrow J/\psi \gamma$	$< 1.4$	$\times 10^{-5}$ CL=90%
$\Gamma_{189} D^*(2010)^- D^+ K^+$	$(6.0 \pm 1.3) \times 10^{-4}$		$\Gamma_{254} X(3930)^0 K^+, X^0 \rightarrow J/\psi \gamma$	$< 2.5$	$\times 10^{-6}$ CL=90%
$\Gamma_{190} D^*(2010)^- D^*(2010)^+ K^+$	$(1.32 \pm 0.18) \times 10^{-3}$		$\Gamma_{255} J/\psi(1S) K^+$	$(1.026 \pm 0.031) \times 10^{-3}$	
$\Gamma_{191} (\bar{D} + \bar{D}^*)(D + D^*) K$	$(4.05 \pm 0.30) \%$		$\Gamma_{256} J/\psi(1S) K^0 \pi^+$		
$\Gamma_{192} D_s^+ \pi^0$	$(1.6 \pm 0.5) \times 10^{-5}$		$\Gamma_{257} J/\psi(1S) K^+ \pi^+ \pi^-$	$(8.1 \pm 1.3) \times 10^{-4}$	S=2.5
$\Gamma_{193} D_s^+ \pi^0$	$< 2.6$	$\times 10^{-4}$ CL=90%	$\Gamma_{258} J/\psi(1S) K^+ K^- K^+$	$(3.37 \pm 0.29) \times 10^{-5}$	
$\Gamma_{194} D_s^+ \eta$	$< 4$	$\times 10^{-4}$ CL=90%	$\Gamma_{259} X(3915) K^+, X \rightarrow p \bar{p}$	$< 7.1$	$\times 10^{-8}$ CL=95%
$\Gamma_{195} D_s^+ \eta$	$< 6$	$\times 10^{-4}$ CL=90%	$\Gamma_{260} J/\psi(1S) K^*(892)^+$	$(1.43 \pm 0.08) \times 10^{-3}$	
$\Gamma_{196} D_s^+ \rho^0$	$< 3.0$	$\times 10^{-4}$ CL=90%	$\Gamma_{261} J/\psi(1S) K(1270)^+$	$(1.8 \pm 0.5) \times 10^{-3}$	
$\Gamma_{197} D_s^+ \rho^0$	$< 4$	$\times 10^{-4}$ CL=90%	$\Gamma_{262} J/\psi(1S) K(1400)^+$	$< 5$	$\times 10^{-4}$ CL=90%
$\Gamma_{198} D_s^+ \omega$	$< 4$	$\times 10^{-4}$ CL=90%	$\Gamma_{263} J/\psi(1S) \eta K^+$	$(1.24 \pm 0.14) \times 10^{-4}$	
$\Gamma_{199} D_s^+ \omega$	$< 6$	$\times 10^{-4}$ CL=90%	$\Gamma_{264} X^{c-odd}(3872) K^+, X^{c-odd} \rightarrow J/\psi \eta$	$< 3.8$	$\times 10^{-6}$ CL=90%
$\Gamma_{200} D_s^+ a_1(1260)^0$	$< 1.8$	$\times 10^{-3}$ CL=90%	$\Gamma_{265} \psi(4160) K^+, \psi \rightarrow J/\psi \eta$	$< 7.4$	$\times 10^{-6}$ CL=90%
$\Gamma_{201} D_s^+ a_1(1260)^0$	$< 1.3$	$\times 10^{-3}$ CL=90%	$\Gamma_{266} J/\psi(1S) \eta' K^+$	$< 8.8$	$\times 10^{-5}$ CL=90%
$\Gamma_{202} D_s^+ \phi$	$(1.7 \pm 1.2) \times 10^{-6}$		$\Gamma_{267} J/\psi(1S) \phi K^+$	$(5.0 \pm 0.4) \times 10^{-5}$	
$\Gamma_{203} D_s^+ \phi$	$< 1.2$	$\times 10^{-5}$ CL=90%	$\Gamma_{268} X(4140) K^+, X \rightarrow J/\psi(1S) \phi$	$(10 \pm 4) \times 10^{-6}$	
$\Gamma_{204} D_s^+ \bar{K}^0$	$< 8$	$\times 10^{-4}$ CL=90%	$\Gamma_{269} X(4274) K^+, X \rightarrow J/\psi(1S) \phi$	$< 4$	$\times 10^{-6}$ CL=90%
$\Gamma_{205} D_s^+ \bar{K}^0$	$< 9$	$\times 10^{-4}$ CL=90%	$\Gamma_{270} J/\psi(1S) \omega K^+$	$(3.20 \pm 0.60) \times 10^{-4}$	
$\Gamma_{206} D_s^+ \bar{K}^*(892)^0$	$< 4.4$	$\times 10^{-6}$ CL=90%	$\Gamma_{271} X(3872) K^+, X \rightarrow J/\psi \omega$	$(6.0 \pm 2.2) \times 10^{-6}$	
$\Gamma_{207} D_s^+ K^{*0}$	$< 3.5$	$\times 10^{-6}$ CL=90%	$\Gamma_{272} X(3915) K^+, X \rightarrow J/\psi \omega$	$(3.0 \pm 0.9) \times 10^{-5}$	
$\Gamma_{208} D_s^+ \bar{K}^*(892)^0$	$< 3.5$	$\times 10^{-4}$ CL=90%	$\Gamma_{273} J/\psi(1S) \pi^+$	$(4.1 \pm 0.4) \times 10^{-5}$	S=2.6
$\Gamma_{209} D_s^+ \pi^+ K^+$	$(1.80 \pm 0.22) \times 10^{-4}$		$\Gamma_{274} J/\psi(1S) \rho^+$	$(5.0 \pm 0.8) \times 10^{-5}$	
$\Gamma_{210} D_s^+ \pi^+ K^+$	$(1.45 \pm 0.24) \times 10^{-4}$		$\Gamma_{275} J/\psi(1S) \pi^+ \pi^0$ nonresonant	$< 7.3$	$\times 10^{-6}$ CL=90%
$\Gamma_{211} D_s^+ \pi^+ K^*(892)^+$	$< 5$	$\times 10^{-3}$ CL=90%	$\Gamma_{276} J/\psi(1S) a_1(1260)^+$	$< 1.2$	$\times 10^{-3}$ CL=90%
$\Gamma_{212} D_s^+ \pi^+ K^*(892)^+$	$< 7$	$\times 10^{-3}$ CL=90%	$\Gamma_{277} J/\psi p \bar{p} \pi^+$	$< 5.0$	$\times 10^{-7}$ CL=90%
$\Gamma_{213} D_s^+ K^+ K^+$	$(9.7 \pm 2.1) \times 10^{-6}$		$\Gamma_{278} J/\psi(1S) p \bar{\Lambda}$	$(1.18 \pm 0.31) \times 10^{-5}$	
$\Gamma_{214} D_s^+ K^+ K^+$	$< 1.5$	$\times 10^{-5}$ CL=90%	$\Gamma_{279} J/\psi(1S) \Sigma^0 p$	$< 1.1$	$\times 10^{-5}$ CL=90%
<b>Charmonium modes</b>			$\Gamma_{280} J/\psi(1S) D^+$	$< 1.2$	$\times 10^{-4}$ CL=90%
$\Gamma_{215} \eta_c K^+$	$(9.6 \pm 1.1) \times 10^{-4}$		$\Gamma_{281} J/\psi(1S) \bar{D}^0 \pi^+$	$< 2.5$	$\times 10^{-5}$ CL=90%
$\Gamma_{216} \eta_c K^+, \eta_c \rightarrow K_S^0 K^\mp \pi^\pm$	$(2.7 \pm 0.6) \times 10^{-5}$		$\Gamma_{282} \psi(2S) \pi^+$	$(2.44 \pm 0.30) \times 10^{-5}$	
$\Gamma_{217} \eta_c K^*(892)^+$	$(1.0 \pm 0.5) \times 10^{-3}$		$\Gamma_{283} \psi(2S) K^+$	$(6.26 \pm 0.24) \times 10^{-4}$	
$\Gamma_{218} \eta_c K^+ \pi^+ \pi^-$	$< 3.9$	$\times 10^{-4}$ CL=90%	$\Gamma_{284} \psi(2S) K^*(892)^+$	$(6.7 \pm 1.4) \times 10^{-4}$	S=1.3
$\Gamma_{219} \eta_c K^+ \omega(782)$	$< 5.3$	$\times 10^{-4}$ CL=90%	$\Gamma_{285} \psi(2S) K^0 \pi^+$		
$\Gamma_{220} \eta_c K^+ \eta$	$< 2.2$	$\times 10^{-4}$ CL=90%	$\Gamma_{286} \psi(2S) K^+ \pi^+ \pi^-$	$(4.3 \pm 0.5) \times 10^{-4}$	
$\Gamma_{221} \eta_c K^+ \pi^0$	$< 6.2$	$\times 10^{-5}$ CL=90%	$\Gamma_{287} \psi(3770) K^+$	$(4.9 \pm 1.3) \times 10^{-4}$	
$\Gamma_{222} \eta_c(2S) K^+$	$(3.4 \pm 1.8) \times 10^{-4}$		$\Gamma_{288} \psi(3770) K^+, \psi \rightarrow D^0 \bar{D}^0$	$(1.5 \pm 0.5) \times 10^{-4}$	S=1.4
$\Gamma_{223} \eta_c(2S) K^+, \eta_c \rightarrow p \bar{p}$	$< 1.06$	$\times 10^{-7}$ CL=95%	$\Gamma_{289} \psi(3770) K^+, \psi \rightarrow D^+ D^-$	$(9.4 \pm 3.5) \times 10^{-5}$	
$\Gamma_{224} \eta_c(2S) K^+, \eta_c \rightarrow K_S^0 K^\mp \pi^\pm$	$(3.4 \pm 2.3) \times 10^{-6}$		$\Gamma_{290} \psi(4040) K^+$	$< 1.3$	$\times 10^{-4}$ CL=90%
$\Gamma_{225} h_c(1P) K^+, h_c \rightarrow J/\psi \pi^+ \pi^-$	$< 3.4$	$\times 10^{-6}$ CL=90%	$\Gamma_{291} \psi(4160) K^+$	$(5.1 \pm 2.7) \times 10^{-4}$	
$\Gamma_{226} X(3730)^0 K^+, X^0 \rightarrow \eta_c \eta$	$< 4.6$	$\times 10^{-5}$ CL=90%	$\Gamma_{292} \psi(4160) K^+, \psi \rightarrow \bar{D}^0 D^0$	$(8 \pm 5) \times 10^{-5}$	
$\Gamma_{227} X(3730)^0 K^+, X^0 \rightarrow \eta_c \pi^0$	$< 5.7$	$\times 10^{-6}$ CL=90%	$\Gamma_{293} \chi_{c0} \pi^+, \chi_{c0} \rightarrow \pi^+ \pi^-$	$< 1$	$\times 10^{-7}$ CL=90%
$\Gamma_{228} X(3872) K^+$	$< 3.2$	$\times 10^{-4}$ CL=90%	$\Gamma_{294} \chi_{c0}(1P) K^+$	$(1.50 \pm 0.15) \times 10^{-4}$	
$\Gamma_{229} X(3872) K^+, X \rightarrow p \bar{p}$	$< 1.7$	$\times 10^{-8}$ CL=95%	$\Gamma_{295} \chi_{c0} K^*(892)^+$	$< 2.1$	$\times 10^{-4}$ CL=90%
$\Gamma_{230} X(3872) K^+, X \rightarrow J/\psi \pi^+ \pi^-$	$(8.6 \pm 0.8) \times 10^{-6}$		$\Gamma_{296} \chi_{c2} \pi^+, \chi_{c2} \rightarrow \pi^+ \pi^-$	$< 1$	$\times 10^{-7}$ CL=90%
$\Gamma_{231} X(3872) K^+, X \rightarrow J/\psi \gamma$	$(2.1 \pm 0.4) \times 10^{-6}$	S=1.1	$\Gamma_{297} \chi_{c2} K^+$	$(1.1 \pm 0.4) \times 10^{-5}$	
$\Gamma_{232} X(3872) K^+, X \rightarrow \psi(2S) \gamma$	$(4 \pm 4) \times 10^{-6}$	S=2.5	$\Gamma_{298} \chi_{c2} K^*(892)^+$	$< 1.2$	$\times 10^{-4}$ CL=90%
$\Gamma_{233} X(3872) K^+, X \rightarrow J/\psi(1S) \eta$	$< 7.7$	$\times 10^{-6}$ CL=90%	$\Gamma_{299} \chi_{c1}(1P) \pi^+$	$(2.2 \pm 0.5) \times 10^{-5}$	
$\Gamma_{234} X(3872) K^+, X \rightarrow D^0 \bar{D}^0$	$< 6.0$	$\times 10^{-5}$ CL=90%	$\Gamma_{300} \chi_{c1}(1P) K^+$	$(4.79 \pm 0.23) \times 10^{-4}$	
$\Gamma_{235} X(3872) K^+, X \rightarrow D^+ D^-$	$< 4.0$	$\times 10^{-5}$ CL=90%	$\Gamma_{301} \chi_{c1}(1P) K^0 \pi^+$		
$\Gamma_{236} X(3872) K^+, X \rightarrow D^0 \bar{D}^0 \pi^0$	$(1.0 \pm 0.4) \times 10^{-4}$		$\Gamma_{302} \chi_{c1}(1P) K^*(892)^+$	$(3.0 \pm 0.6) \times 10^{-4}$	S=1.1
$\Gamma_{237} X(3872) K^+, X \rightarrow \bar{D}^{*0} D^0$	$(8.5 \pm 2.6) \times 10^{-5}$	S=1.4	$\Gamma_{303} h_c(1P) K^+$	$< 3.8$	$\times 10^{-5}$ CL=90%
$\Gamma_{238} X(3872)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-$	$< 3.0$	$\times 10^{-5}$ CL=90%	$\Gamma_{304} h_c(1P) K^+, h_c \rightarrow p \bar{p}$	$< 6.4$	$\times 10^{-8}$ CL=95%
$\Gamma_{239} X(3872)^0 K^+, X^0 \rightarrow \eta_c \omega(782)$	$< 6.9$	$\times 10^{-5}$ CL=90%			
$\Gamma_{240} X(3915)^0 K^+, X^0 \rightarrow \eta_c \eta$	$< 3.3$	$\times 10^{-5}$ CL=90%	<b>K or K* modes</b>		
$\Gamma_{241} X(3915)^0 K^+, X^0 \rightarrow \eta_c \pi^0$	$< 1.8$	$\times 10^{-5}$ CL=90%	$\Gamma_{305} K^0 \pi^+$	$(2.37 \pm 0.08) \times 10^{-5}$	
$\Gamma_{242} X(4014)^0 K^+, X^0 \rightarrow \eta_c \eta$	$< 3.9$	$\times 10^{-5}$ CL=90%	$\Gamma_{306} K^+ \pi^0$	$(1.29 \pm 0.05) \times 10^{-5}$	
$\Gamma_{243} X(4014)^0 K^+, X^0 \rightarrow \eta_c \pi^0$	$< 1.2$	$\times 10^{-5}$ CL=90%	$\Gamma_{307} \eta' K^+$	$(7.06 \pm 0.25) \times 10^{-5}$	
$\Gamma_{244} X(3900)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-$	$< 4.7$	$\times 10^{-5}$ CL=90%	$\Gamma_{308} \eta' K^*(892)^+$	$(4.8 \pm 1.8) \times 10^{-6}$	
$\Gamma_{245} X(4020)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-$	$< 1.6$	$\times 10^{-5}$ CL=90%			

## Meson Particle Listings

 $B^\pm$ 

$\Gamma_{309}$	$\eta' K_0^*(1430)^+$	$(5.2 \pm 2.1) \times 10^{-6}$	
$\Gamma_{310}$	$\eta' K_2^*(1430)^+$	$(2.8 \pm 0.5) \times 10^{-5}$	
$\Gamma_{311}$	$\eta K^+$	$(2.4 \pm 0.4) \times 10^{-6}$	S=1.7
$\Gamma_{312}$	$\eta K^*(892)^+$	$(1.93 \pm 0.16) \times 10^{-5}$	
$\Gamma_{313}$	$\eta K_0^*(1430)^+$	$(1.8 \pm 0.4) \times 10^{-5}$	
$\Gamma_{314}$	$\eta K_2^*(1430)^+$	$(9.1 \pm 3.0) \times 10^{-6}$	
$\Gamma_{315}$	$\eta(1295) K^+ \times B(\eta(1295) \rightarrow \eta \pi \pi)$	$(2.9 \pm 0.8) \times 10^{-6}$	
$\Gamma_{316}$	$\eta(1405) K^+ \times B(\eta(1405) \rightarrow \eta \pi \pi)$	$< 1.3$	$\times 10^{-6}$ CL=90%
$\Gamma_{317}$	$\eta(1405) K^+ \times B(\eta(1405) \rightarrow K^* K)$	$< 1.2$	$\times 10^{-6}$ CL=90%
$\Gamma_{318}$	$\eta(1475) K^+ \times B(\eta(1475) \rightarrow K^* K)$	$(1.38 \pm 0.21) \times 10^{-5}$	
$\Gamma_{319}$	$f_1(1285) K^+$	$< 2.0$	$\times 10^{-6}$ CL=90%
$\Gamma_{320}$	$f_1(1420) K^+ \times B(f_1(1420) \rightarrow \eta \pi \pi)$	$< 2.9$	$\times 10^{-6}$ CL=90%
$\Gamma_{321}$	$f_1(1420) K^+ \times B(f_1(1420) \rightarrow K^* K)$	$< 4.1$	$\times 10^{-6}$ CL=90%
$\Gamma_{322}$	$\phi(1680) K^+ \times B(\phi(1680) \rightarrow K^* K)$	$< 3.4$	$\times 10^{-6}$ CL=90%
$\Gamma_{323}$	$f_0(1500) K^+$	$(3.7 \pm 2.2) \times 10^{-6}$	
$\Gamma_{324}$	$\omega K^+$	$(6.5 \pm 0.4) \times 10^{-6}$	
$\Gamma_{325}$	$\omega K^*(892)^+$	$< 7.4$	$\times 10^{-6}$ CL=90%
$\Gamma_{326}$	$\omega(K\pi)_0^{*+}$	$(2.8 \pm 0.4) \times 10^{-5}$	
$\Gamma_{327}$	$\omega K_0^*(1430)^+$	$(2.4 \pm 0.5) \times 10^{-5}$	
$\Gamma_{328}$	$\omega K_2^*(1430)^+$	$(2.1 \pm 0.4) \times 10^{-5}$	
$\Gamma_{329}$	$a_0(980)^+ K^0 \times B(a_0(980)^+ \rightarrow \eta \pi^+)$	$< 3.9$	$\times 10^{-6}$ CL=90%
$\Gamma_{330}$	$a_0(980)^0 K^+ \times B(a_0(980)^0 \rightarrow \eta \pi^0)$	$< 2.5$	$\times 10^{-6}$ CL=90%
$\Gamma_{331}$	$K^*(892)^0 \pi^+$	$(1.01 \pm 0.09) \times 10^{-5}$	
$\Gamma_{332}$	$K^*(892)^+ \pi^0$	$(8.2 \pm 1.9) \times 10^{-6}$	
$\Gamma_{333}$	$K^+ \pi^- \pi^+$	$(5.10 \pm 0.29) \times 10^{-5}$	
$\Gamma_{334}$	$K^+ \pi^- \pi^+$ nonresonant	$(1.63 \pm 0.21) \times 10^{-5}$	
$\Gamma_{335}$	$\omega(782) K^+$	$(6 \pm 9) \times 10^{-6}$	
$\Gamma_{336}$	$K^+ f_0(980) \times B(f_0(980) \rightarrow \pi^+ \pi^-)$	$(9.4 \pm 1.0) \times 10^{-6}$	
$\Gamma_{337}$	$f_2(1270)^0 K^+$	$(1.07 \pm 0.27) \times 10^{-6}$	
$\Gamma_{338}$	$f_0(1370)^0 K^+ \times B(f_0(1370)^0 \rightarrow \pi^+ \pi^-)$	$< 1.07$	$\times 10^{-5}$ CL=90%
$\Gamma_{339}$	$\rho^0(1450) K^+ \times B(\rho^0(1450) \rightarrow \pi^+ \pi^-)$	$< 1.17$	$\times 10^{-5}$ CL=90%
$\Gamma_{340}$	$f_2'(1525) K^+ \times B(f_2'(1525) \rightarrow \pi^+ \pi^-)$	$< 3.4$	$\times 10^{-6}$ CL=90%
$\Gamma_{341}$	$K^+ \rho^0$	$(3.7 \pm 0.5) \times 10^{-6}$	
$\Gamma_{342}$	$K_0^*(1430)^0 \pi^+$	$(4.5 \pm 0.9) \times 10^{-5}$	S=1.5
$\Gamma_{343}$	$K_2^*(1430)^0 \pi^+$	$(5.6 \pm 2.2) \times 10^{-6}$	
$\Gamma_{344}$	$K^*(1410)^0 \pi^+$	$< 4.5$	$\times 10^{-5}$ CL=90%
$\Gamma_{345}$	$K^*(1680)^0 \pi^+$	$< 1.2$	$\times 10^{-5}$ CL=90%
$\Gamma_{346}$	$K^+ \pi^0 \pi^0$	$(1.62 \pm 0.19) \times 10^{-5}$	
$\Gamma_{347}$	$f_0(980) K^+ \times B(f_0 \rightarrow \pi^0 \pi^0)$	$(2.8 \pm 0.8) \times 10^{-6}$	
$\Gamma_{348}$	$K^- \pi^+ \pi^+$	$< 9.5$	$\times 10^{-7}$ CL=90%
$\Gamma_{349}$	$K^- \pi^+ \pi^+$ nonresonant	$< 5.6$	$\times 10^{-5}$ CL=90%
$\Gamma_{350}$	$K_1(1270)^0 \pi^+$	$< 4.0$	$\times 10^{-5}$ CL=90%
$\Gamma_{351}$	$K_1(1400)^0 \pi^+$	$< 3.9$	$\times 10^{-5}$ CL=90%
$\Gamma_{352}$	$K^0 \pi^+ \pi^0$	$< 6.6$	$\times 10^{-5}$ CL=90%
$\Gamma_{353}$	$K^0 \rho^+$	$(8.0 \pm 1.5) \times 10^{-6}$	
$\Gamma_{354}$	$K^*(892)^+ \pi^+ \pi^-$	$(7.5 \pm 1.0) \times 10^{-5}$	
$\Gamma_{355}$	$K^*(892)^+ \rho^0$	$(4.6 \pm 1.1) \times 10^{-6}$	
$\Gamma_{356}$	$K^*(892)^+ f_0(980)$	$(4.2 \pm 0.7) \times 10^{-6}$	
$\Gamma_{357}$	$a_1^+ K^0$	$(3.5 \pm 0.7) \times 10^{-5}$	
$\Gamma_{358}$	$b_1^+ K^0 \times B(b_1^+ \rightarrow \omega \pi^+)$	$(9.6 \pm 1.9) \times 10^{-6}$	
$\Gamma_{359}$	$K^*(892)^0 \rho^+$	$(9.2 \pm 1.5) \times 10^{-6}$	
$\Gamma_{360}$	$K_1(1400)^+ \rho^0$	$< 7.8$	$\times 10^{-4}$ CL=90%
$\Gamma_{361}$	$K_2^*(1430)^+ \rho^0$	$< 1.5$	$\times 10^{-3}$ CL=90%
$\Gamma_{362}$	$b_1^0 K^+ \times B(b_1^0 \rightarrow \omega \pi^0)$	$(9.1 \pm 2.0) \times 10^{-6}$	
$\Gamma_{363}$	$b_1^+ K^* \times B(b_1^+ \rightarrow \omega \pi^+)$	$< 5.9$	$\times 10^{-6}$ CL=90%
$\Gamma_{364}$	$b_1^0 K^* \times B(b_1^0 \rightarrow \omega \pi^0)$	$< 6.7$	$\times 10^{-6}$ CL=90%
$\Gamma_{365}$	$K^+ \bar{K}^0$	$(1.31 \pm 0.17) \times 10^{-6}$	S=1.2
$\Gamma_{366}$	$\bar{K}^0 K^+ \pi^0$	$< 2.4$	$\times 10^{-5}$ CL=90%

$\Gamma_{367}$	$K^+ K_S^0 K_S^0$	$(1.08 \pm 0.06) \times 10^{-5}$	
$\Gamma_{368}$	$f_0(980) K^+, f_0 \rightarrow K_S^0 K_S^0$	$(1.47 \pm 0.33) \times 10^{-5}$	
$\Gamma_{369}$	$f_0(1710) K^+, f_0 \rightarrow K_S^0 K_S^0$	$(4.8 \pm 4.0) \times 10^{-7}$	
$\Gamma_{370}$	$K^+ K_S^0 K_S^0$ nonresonant	$(2.0 \pm 0.4) \times 10^{-5}$	
$\Gamma_{371}$	$K_S^0 K_S^0 \pi^+$	$< 5.1$	$\times 10^{-7}$ CL=90%
$\Gamma_{372}$	$K^+ K^- \pi^+$	$(5.0 \pm 0.7) \times 10^{-6}$	
$\Gamma_{373}$	$K^+ K^- \pi^+$ nonresonant	$< 7.5$	$\times 10^{-5}$ CL=90%
$\Gamma_{374}$	$K^+ \bar{K}^*(892)^0$	$< 1.1$	$\times 10^{-6}$ CL=90%
$\Gamma_{375}$	$K^+ \bar{K}_0^*(1430)^0$	$< 2.2$	$\times 10^{-6}$ CL=90%
$\Gamma_{376}$	$K^+ K^+ \pi^-$	$< 1.6$	$\times 10^{-7}$ CL=90%
$\Gamma_{377}$	$K^+ K^+ \pi^-$ nonresonant	$< 8.79$	$\times 10^{-5}$ CL=90%
$\Gamma_{378}$	$f_2'(1525) K^+$	$(1.8 \pm 0.5) \times 10^{-6}$	S=1.1
$\Gamma_{379}$	$K^+ f_J(2220)$		
$\Gamma_{380}$	$K^* \pi^+ K^-$	$< 1.18$	$\times 10^{-5}$ CL=90%
$\Gamma_{381}$	$K^*(892)^+ K^*(892)^0$	$(9.1 \pm 2.9) \times 10^{-7}$	
$\Gamma_{382}$	$K^* \pi^+ K^-$	$< 6.1$	$\times 10^{-6}$ CL=90%
$\Gamma_{383}$	$K^+ K^- K^+$	$(3.40 \pm 0.14) \times 10^{-5}$	S=1.4
$\Gamma_{384}$	$K^+ \phi$	$(8.8 \pm 0.7) \times 10^{-6}$	S=1.1
$\Gamma_{385}$	$f_0(980) K^+ \times B(f_0(980) \rightarrow K^+ K^-)$	$(9.4 \pm 3.2) \times 10^{-6}$	
$\Gamma_{386}$	$a_2(1320) K^+ \times B(a_2(1320) \rightarrow K^+ K^-)$	$< 1.1$	$\times 10^{-6}$ CL=90%
$\Gamma_{387}$	$X_0(1550) K^+ \times B(X_0(1550) \rightarrow K^+ K^-)$	$(4.3 \pm 0.7) \times 10^{-6}$	
$\Gamma_{388}$	$\phi(1680) K^+ \times B(\phi(1680) \rightarrow K^+ K^-)$	$< 8$	$\times 10^{-7}$ CL=90%
$\Gamma_{389}$	$f_0(1710) K^+ \times B(f_0(1710) \rightarrow K^+ K^-)$	$(1.1 \pm 0.6) \times 10^{-6}$	
$\Gamma_{390}$	$K^+ K^- K^+$ nonresonant	$(2.38 \pm 0.28) \times 10^{-5}$	
$\Gamma_{391}$	$K^*(892)^+ K^+ K^-$	$(3.6 \pm 0.5) \times 10^{-5}$	
$\Gamma_{392}$	$K^*(892)^+ \phi$	$(10.0 \pm 2.0) \times 10^{-6}$	S=1.7
$\Gamma_{393}$	$\phi(K\pi)_0^{*+}$	$(8.3 \pm 1.6) \times 10^{-6}$	
$\Gamma_{394}$	$\phi K_1(1270)^+$	$(6.1 \pm 1.9) \times 10^{-6}$	
$\Gamma_{395}$	$\phi K_1(1400)^+$	$< 3.2$	$\times 10^{-6}$ CL=90%
$\Gamma_{396}$	$\phi K^*(1410)^+$	$< 4.3$	$\times 10^{-6}$ CL=90%
$\Gamma_{397}$	$\phi K_0^*(1430)^+$	$(7.0 \pm 1.6) \times 10^{-6}$	
$\Gamma_{398}$	$\phi K_2^*(1430)^+$	$(8.4 \pm 2.1) \times 10^{-6}$	
$\Gamma_{399}$	$\phi K_2^*(1770)^+$	$< 1.50$	$\times 10^{-5}$ CL=90%
$\Gamma_{400}$	$\phi K_2^*(1820)^+$	$< 1.63$	$\times 10^{-5}$ CL=90%
$\Gamma_{401}$	$a_1^+ K^* \pi^0$	$< 3.6$	$\times 10^{-6}$ CL=90%
$\Gamma_{402}$	$K^+ \phi \phi$	$(5.0 \pm 1.2) \times 10^{-6}$	S=2.3
$\Gamma_{403}$	$\eta' \eta' K^+$	$< 2.5$	$\times 10^{-5}$ CL=90%
$\Gamma_{404}$	$\omega \phi K^+$	$< 1.9$	$\times 10^{-6}$ CL=90%
$\Gamma_{405}$	$X(1812) K^+ \times B(X \rightarrow \omega \phi)$	$< 3.2$	$\times 10^{-7}$ CL=90%
$\Gamma_{406}$	$K^*(892)^+ \gamma$	$(4.21 \pm 0.18) \times 10^{-5}$	
$\Gamma_{407}$	$K_1(1270)^+ \gamma$	$(4.3 \pm 1.3) \times 10^{-5}$	
$\Gamma_{408}$	$\eta K^+ \gamma$	$(7.9 \pm 0.9) \times 10^{-6}$	
$\Gamma_{409}$	$\eta' K^+ \gamma$	$(2.9 \pm 1.0) \times 10^{-6}$	
$\Gamma_{410}$	$\phi K^+ \gamma$	$(2.7 \pm 0.4) \times 10^{-6}$	S=1.2
$\Gamma_{411}$	$K^+ \pi^- \pi^+ \gamma$	$(2.76 \pm 0.22) \times 10^{-5}$	S=1.2
$\Gamma_{412}$	$K^*(892)^0 \pi^+ \gamma$	$(2.0 \pm 0.7) \times 10^{-5}$	
$\Gamma_{413}$	$K^+ \rho^0 \gamma$	$< 2.0$	$\times 10^{-5}$ CL=90%
$\Gamma_{414}$	$K^+ \pi^- \pi^+ \gamma$ nonresonant	$< 9.2$	$\times 10^{-6}$ CL=90%
$\Gamma_{415}$	$K^0 \pi^+ \pi^0 \gamma$	$(4.6 \pm 0.5) \times 10^{-5}$	
$\Gamma_{416}$	$K_1(1400)^+ \gamma$	$< 1.5$	$\times 10^{-5}$ CL=90%
$\Gamma_{417}$	$K_2^*(1430)^+ \gamma$	$(1.4 \pm 0.4) \times 10^{-5}$	
$\Gamma_{418}$	$K^*(1680)^+ \gamma$	$< 1.9$	$\times 10^{-3}$ CL=90%
$\Gamma_{419}$	$K_2^*(1780)^+ \gamma$	$< 3.9$	$\times 10^{-5}$ CL=90%
$\Gamma_{420}$	$K_4^*(2045)^+ \gamma$	$< 9.9$	$\times 10^{-3}$ CL=90%

## Light unflavored meson modes

$\Gamma_{421}$	$\rho^+ \gamma$	$(9.8 \pm 2.5) \times 10^{-7}$	
$\Gamma_{422}$	$\pi^+ \pi^0$	$(5.5 \pm 0.4) \times 10^{-6}$	S=1.2
$\Gamma_{423}$	$\pi^+ \pi^+ \pi^-$	$(1.52 \pm 0.14) \times 10^{-5}$	
$\Gamma_{424}$	$\rho^0 \pi^+$	$(8.3 \pm 1.2) \times 10^{-6}$	
$\Gamma_{425}$	$\pi^+ f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 1.5$	$\times 10^{-6}$ CL=90%
$\Gamma_{426}$	$\pi^+ f_2(1270)$	$(1.6 \pm 0.7) \times 10^{-6}$	
$\Gamma_{427}$	$\rho(1450)^0 \pi^+, \rho^0 \rightarrow \pi^+ \pi^-$	$(1.4 \pm 0.6) \times 10^{-6}$	
$\Gamma_{428}$	$f_0(1370) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	$< 4.0$	$\times 10^{-6}$ CL=90%
$\Gamma_{429}$	$f_0(500) \pi^+, f_0 \rightarrow \pi^+ \pi^-$	$< 4.1$	$\times 10^{-6}$ CL=90%

$\Gamma_{430}$	$\pi^+\pi^-\pi^+$ nonresonant	( 5.3 $\pm$ 1.5 ) $\times 10^{-6}$	
$\Gamma_{431}$	$\pi^+\pi^0\pi^0$	< 8.9 $\times 10^{-4}$	CL=90%
$\Gamma_{432}$	$\rho^+\pi^0$	( 1.09 $\pm$ 0.14 ) $\times 10^{-5}$	
$\Gamma_{433}$	$\pi^+\pi^-\pi^+\pi^0$	< 4.0 $\times 10^{-3}$	CL=90%
$\Gamma_{434}$	$\rho^+\rho^0$	( 2.40 $\pm$ 0.19 ) $\times 10^{-5}$	
$\Gamma_{435}$	$\rho^+f_0(980), f_0 \rightarrow \pi^+\pi^-$	< 2.0 $\times 10^{-6}$	CL=90%
$\Gamma_{436}$	$a_1(1260)^+\pi^0$	( 2.6 $\pm$ 0.7 ) $\times 10^{-5}$	
$\Gamma_{437}$	$a_1(1260)^0\pi^+$	( 2.0 $\pm$ 0.6 ) $\times 10^{-5}$	
$\Gamma_{438}$	$\omega\pi^+$	( 6.9 $\pm$ 0.5 ) $\times 10^{-6}$	
$\Gamma_{439}$	$\omega\rho^+$	( 1.59 $\pm$ 0.21 ) $\times 10^{-5}$	
$\Gamma_{440}$	$\eta\pi^+$	( 4.02 $\pm$ 0.27 ) $\times 10^{-6}$	
$\Gamma_{441}$	$\eta\rho^+$	( 7.0 $\pm$ 2.9 ) $\times 10^{-6}$	S=2.8
$\Gamma_{442}$	$\eta'\pi^+$	( 2.7 $\pm$ 0.9 ) $\times 10^{-6}$	S=1.9
$\Gamma_{443}$	$\eta'\rho^+$	( 9.7 $\pm$ 2.2 ) $\times 10^{-6}$	
$\Gamma_{444}$	$\phi\pi^+$	< 1.5 $\times 10^{-7}$	CL=90%
$\Gamma_{445}$	$\phi\rho^+$	< 3.0 $\times 10^{-6}$	CL=90%
$\Gamma_{446}$	$a_0(980)^0\pi^+, a_0^0 \rightarrow \eta\pi^0$	< 5.8 $\times 10^{-6}$	CL=90%
$\Gamma_{447}$	$a_0(980)^+\pi^0, a_0^+ \rightarrow \eta\pi^+$	< 1.4 $\times 10^{-6}$	CL=90%
$\Gamma_{448}$	$\pi^+\pi^+\pi^+\pi^-\pi^-$	< 8.6 $\times 10^{-4}$	CL=90%
$\Gamma_{449}$	$\rho^0 a_1(1260)^+$	< 6.2 $\times 10^{-4}$	CL=90%
$\Gamma_{450}$	$\rho^0 a_2(1320)^+$	< 7.2 $\times 10^{-4}$	CL=90%
$\Gamma_{451}$	$b_1^0\pi^+, b_1^0 \rightarrow \omega\pi^0$	( 6.7 $\pm$ 2.0 ) $\times 10^{-6}$	
$\Gamma_{452}$	$b_1^+\pi^0, b_1^+ \rightarrow \omega\pi^+$	< 3.3 $\times 10^{-6}$	CL=90%
$\Gamma_{453}$	$\pi^+\pi^+\pi^+\pi^-\pi^-$	< 6.3 $\times 10^{-3}$	CL=90%
$\Gamma_{454}$	$b_1^+\rho^0, b_1^+ \rightarrow \omega\pi^+$	< 5.2 $\times 10^{-6}$	CL=90%
$\Gamma_{455}$	$a_1(1260)^+a_1(1260)^0$	< 1.3 $\times 10^{-6}$	CL=90%
$\Gamma_{456}$	$b_1^0\rho^+, b_1^0 \rightarrow \omega\pi^0$	< 3.3 $\times 10^{-6}$	CL=90%

Charged particle ( $h^\pm$ ) modes

	$h^\pm = K^\pm$ or $\pi^\pm$		
$\Gamma_{457}$	$h^+\pi^0$	( 1.6 $\pm$ 0.7 ) $\times 10^{-5}$	
$\Gamma_{458}$	$\omega h^+$	( 1.38 $\pm$ 0.27 ) $\times 10^{-5}$	
$\Gamma_{459}$	$h^+X^0$ (Familon)	< 4.9 $\times 10^{-5}$	CL=90%

## Baryon modes

$\Gamma_{460}$	$p\bar{p}\pi^+$	( 1.62 $\pm$ 0.20 ) $\times 10^{-6}$	
$\Gamma_{461}$	$p\bar{p}\pi^+$ nonresonant	< 5.3 $\times 10^{-5}$	CL=90%
$\Gamma_{462}$	$p\bar{p}\pi^+\pi^+\pi^-$		
$\Gamma_{463}$	$p\bar{p}K^+$	( 5.9 $\pm$ 0.5 ) $\times 10^{-6}$	S=1.5
$\Gamma_{464}$	$\Theta(1710)^{++}\bar{p}, \Theta^{++} \rightarrow pK^+$	[g] < 9.1 $\times 10^{-8}$	CL=90%
$\Gamma_{465}$	$f_J(2220)K^+, f_J \rightarrow p\bar{p}$	[g] < 4.1 $\times 10^{-7}$	CL=90%
$\Gamma_{466}$	$p\bar{\Lambda}(1520)$	( 3.1 $\pm$ 0.6 ) $\times 10^{-7}$	
$\Gamma_{467}$	$p\bar{p}K^+$ nonresonant	< 8.9 $\times 10^{-5}$	CL=90%
$\Gamma_{468}$	$p\bar{p}K^*(892)^+$	( 3.6 $\pm$ 0.8 ) $\times 10^{-6}$	
$\Gamma_{469}$	$f_J(2220)K^{*+}, f_J \rightarrow p\bar{p}$	< 7.7 $\times 10^{-7}$	CL=90%
$\Gamma_{470}$	$p\bar{\Lambda}$	< 3.2 $\times 10^{-7}$	CL=90%
$\Gamma_{471}$	$p\bar{\Lambda}\gamma$	( 2.4 $\pm$ 0.5 ) $\times 10^{-6}$	
$\Gamma_{472}$	$p\bar{\Lambda}\pi^0$	( 3.0 $\pm$ 0.7 ) $\times 10^{-6}$	
$\Gamma_{473}$	$\rho\bar{\Sigma}(1385)^0$	< 4.7 $\times 10^{-7}$	CL=90%
$\Gamma_{474}$	$\Delta^+\bar{\Lambda}$	< 8.2 $\times 10^{-7}$	CL=90%
$\Gamma_{475}$	$\rho\bar{\Sigma}\gamma$	< 4.6 $\times 10^{-6}$	CL=90%
$\Gamma_{476}$	$\rho\bar{\Lambda}\pi^+\pi^-$	( 5.9 $\pm$ 1.1 ) $\times 10^{-6}$	
$\Gamma_{477}$	$\rho\bar{\Lambda}\rho^0$	( 4.8 $\pm$ 0.9 ) $\times 10^{-6}$	
$\Gamma_{478}$	$\rho\bar{\Lambda}f_2(1270)$	( 2.0 $\pm$ 0.8 ) $\times 10^{-6}$	
$\Gamma_{479}$	$\Lambda\bar{\Lambda}\pi^+$	< 9.4 $\times 10^{-7}$	CL=90%
$\Gamma_{480}$	$\Lambda\bar{\Lambda}K^+$	( 3.4 $\pm$ 0.6 ) $\times 10^{-6}$	
$\Gamma_{481}$	$\Lambda\bar{\Lambda}K^{*+}$	( 2.2 $\pm$ 1.2 ) $\times 10^{-6}$	
$\Gamma_{482}$	$\bar{\Delta}^0 p$	< 1.38 $\times 10^{-6}$	CL=90%
$\Gamma_{483}$	$\Delta^{++}\bar{p}$	< 1.4 $\times 10^{-7}$	CL=90%
$\Gamma_{484}$	$D^+\bar{p}\bar{p}$	< 1.5 $\times 10^{-5}$	CL=90%
$\Gamma_{485}$	$D^*(2010)^+\bar{p}\bar{p}$	< 1.5 $\times 10^{-5}$	CL=90%
$\Gamma_{486}$	$\bar{D}^0 p\bar{p}\pi^+$	( 3.72 $\pm$ 0.27 ) $\times 10^{-4}$	
$\Gamma_{487}$	$\bar{D}^{*0} p\bar{p}\pi^+$	( 3.73 $\pm$ 0.32 ) $\times 10^{-4}$	
$\Gamma_{488}$	$D^-\bar{p}\bar{p}\pi^+\pi^-$	( 1.66 $\pm$ 0.30 ) $\times 10^{-4}$	
$\Gamma_{489}$	$D^{*-}\bar{p}\bar{p}\pi^+\pi^-$	( 1.86 $\pm$ 0.25 ) $\times 10^{-4}$	
$\Gamma_{490}$	$\rho\bar{\Lambda}^0\bar{D}^0$	( 1.43 $\pm$ 0.32 ) $\times 10^{-5}$	
$\Gamma_{491}$	$\rho\bar{\Lambda}^0\bar{D}^*(2007)^0$	< 5 $\times 10^{-5}$	CL=90%
$\Gamma_{492}$	$\bar{\Lambda}_c^- p\pi^+$	( 2.2 $\pm$ 0.4 ) $\times 10^{-4}$	S=2.2
$\Gamma_{493}$	$\bar{\Lambda}_c^- \Delta(1232)^{++}$	< 1.9 $\times 10^{-5}$	CL=90%

$\Gamma_{494}$	$\bar{\Lambda}_c^- \Delta_X(1600)^{++}$	( 4.6 $\pm$ 0.9 ) $\times 10^{-5}$	
$\Gamma_{495}$	$\bar{\Lambda}_c^- \Delta_X(2420)^{++}$	( 3.7 $\pm$ 0.8 ) $\times 10^{-5}$	
$\Gamma_{496}$	$(\bar{\Lambda}_c^- p)_s \pi^+$	[h] ( 3.1 $\pm$ 0.7 ) $\times 10^{-5}$	
$\Gamma_{497}$	$\bar{\Sigma}_c(2520)^0 p$	< 3 $\times 10^{-6}$	CL=90%
$\Gamma_{498}$	$\bar{\Sigma}_c(2800)^0 p$	( 2.6 $\pm$ 0.9 ) $\times 10^{-5}$	
$\Gamma_{499}$	$\bar{\Lambda}_c^- p\pi^+\pi^0$	( 1.8 $\pm$ 0.6 ) $\times 10^{-3}$	
$\Gamma_{500}$	$\bar{\Lambda}_c^- p\pi^+\pi^+\pi^-$	( 2.2 $\pm$ 0.7 ) $\times 10^{-3}$	
$\Gamma_{501}$	$\bar{\Lambda}_c^- p\pi^+\pi^+\pi^-\pi^0$	< 1.34 %	CL=90%
$\Gamma_{502}$	$\Lambda_c^+ \Lambda_c^- K^+$	( 6.9 $\pm$ 2.2 ) $\times 10^{-4}$	
$\Gamma_{503}$	$\bar{\Sigma}_c(2455)^0 p$	( 2.9 $\pm$ 0.7 ) $\times 10^{-5}$	
$\Gamma_{504}$	$\bar{\Sigma}_c(2455)^0 p\pi^0$	( 3.5 $\pm$ 1.1 ) $\times 10^{-4}$	
$\Gamma_{505}$	$\bar{\Sigma}_c(2455)^0 p\pi^-\pi^+$	( 3.5 $\pm$ 1.0 ) $\times 10^{-4}$	
$\Gamma_{506}$	$\bar{\Sigma}_c(2455)^{--} p\pi^+\pi^+$	( 2.34 $\pm$ 0.20 ) $\times 10^{-4}$	
$\Gamma_{507}$	$\bar{\Lambda}_c(2593)^- / \bar{\Lambda}_c(2625)^- p\pi^+$	< 1.9 $\times 10^{-4}$	CL=90%
$\Gamma_{508}$	$\Xi_c^0 \Lambda_c^+, \Xi_c^0 \rightarrow \Xi^+ \pi^-$	( 2.4 $\pm$ 0.9 ) $\times 10^{-5}$	S=1.4
$\Gamma_{509}$	$\Xi_c^0 \Lambda_c^+, \Xi_c^0 \rightarrow \Lambda K^+ \pi^-$	( 2.1 $\pm$ 0.9 ) $\times 10^{-5}$	S=1.5

Lepton Family number ( $LF$ ) or Lepton number ( $L$ ) or Baryon number ( $B$ ) violating modes, or/and  $\Delta B = 1$  weak neutral current ( $B1$ ) modes

$\Gamma_{510}$	$\pi^+ \ell^+ \ell^-$	$B1$	< 4.9 $\times 10^{-8}$	CL=90%
$\Gamma_{511}$	$\pi^+ e^+ e^-$	$B1$	< 8.0 $\times 10^{-8}$	CL=90%
$\Gamma_{512}$	$\pi^+ \mu^+ \mu^-$	$B1$	( 1.79 $\pm$ 0.23 ) $\times 10^{-8}$	
$\Gamma_{513}$	$\pi^+ \nu \bar{\nu}$	$B1$	< 9.8 $\times 10^{-5}$	CL=90%
$\Gamma_{514}$	$K^+ \ell^+ \ell^-$	$B1$	[a] ( 4.51 $\pm$ 0.23 ) $\times 10^{-7}$	S=1.1
$\Gamma_{515}$	$K^+ e^+ e^-$	$B1$	( 5.5 $\pm$ 0.7 ) $\times 10^{-7}$	
$\Gamma_{516}$	$K^+ \mu^+ \mu^-$	$B1$	( 4.43 $\pm$ 0.24 ) $\times 10^{-7}$	S=1.2
$\Gamma_{517}$	$K^+ \nu \bar{\nu}$	$B1$	< 1.6 $\times 10^{-5}$	CL=90%
$\Gamma_{518}$	$\rho^+ \nu \bar{\nu}$	$B1$	< 2.13 $\times 10^{-4}$	CL=90%
$\Gamma_{519}$	$K^*(892)^+ \ell^+ \ell^-$	$B1$	[a] ( 1.01 $\pm$ 0.11 ) $\times 10^{-6}$	S=1.1
$\Gamma_{520}$	$K^*(892)^+ e^+ e^-$	$B1$	( 1.55 $\pm$ 0.40 ) $\times 10^{-6}$	
$\Gamma_{521}$	$K^*(892)^+ \mu^+ \mu^-$	$B1$	( 9.6 $\pm$ 1.0 ) $\times 10^{-7}$	
$\Gamma_{522}$	$K^*(892)^+ \nu \bar{\nu}$	$B1$	< 4.0 $\times 10^{-5}$	CL=90%
$\Gamma_{523}$	$K^+ \pi^+ \pi^- \mu^+ \mu^-$	$B1$	( 4.4 $\pm$ 0.4 ) $\times 10^{-7}$	
$\Gamma_{524}$	$\phi K^+ \mu^+ \mu^-$	$B1$	( 7.9 $\pm$ 2.1 ) $\times 10^{-8}$	
$\Gamma_{525}$	$\pi^+ e^+ \mu^-$	$LF$	< 6.4 $\times 10^{-3}$	CL=90%
$\Gamma_{526}$	$\pi^+ e^- \mu^+$	$LF$	< 6.4 $\times 10^{-3}$	CL=90%
$\Gamma_{527}$	$\pi^+ e^\pm \mu^\mp$	$LF$	< 1.7 $\times 10^{-7}$	CL=90%
$\Gamma_{528}$	$\pi^+ e^+ \tau^-$	$LF$	< 7.4 $\times 10^{-5}$	CL=90%
$\Gamma_{529}$	$\pi^+ e^- \tau^+$	$LF$	< 2.0 $\times 10^{-5}$	CL=90%
$\Gamma_{530}$	$\pi^+ e^\pm \tau^\mp$	$LF$	< 7.5 $\times 10^{-5}$	CL=90%
$\Gamma_{531}$	$\pi^+ \mu^+ \tau^-$	$LF$	< 6.2 $\times 10^{-5}$	CL=90%
$\Gamma_{532}$	$\pi^+ \mu^- \tau^+$	$LF$	< 4.5 $\times 10^{-5}$	CL=90%
$\Gamma_{533}$	$\pi^+ \mu^\pm \tau^\mp$	$LF$	< 7.2 $\times 10^{-5}$	CL=90%
$\Gamma_{534}$	$K^+ e^+ \mu^-$	$LF$	< 9.1 $\times 10^{-8}$	CL=90%
$\Gamma_{535}$	$K^+ e^- \mu^+$	$LF$	< 1.3 $\times 10^{-7}$	CL=90%
$\Gamma_{536}$	$K^+ e^\pm \mu^\mp$	$LF$	< 9.1 $\times 10^{-8}$	CL=90%
$\Gamma_{537}$	$K^+ e^+ \tau^-$	$LF$	< 4.3 $\times 10^{-5}$	CL=90%
$\Gamma_{538}$	$K^+ e^- \tau^+$	$LF$	< 1.5 $\times 10^{-5}$	CL=90%
$\Gamma_{539}$	$K^+ e^\pm \tau^\mp$	$LF$	< 3.0 $\times 10^{-5}$	CL=90%
$\Gamma_{540}$	$K^+ \mu^+ \tau^-$	$LF$	< 4.5 $\times 10^{-5}$	CL=90%
$\Gamma_{541}$	$K^+ \mu^- \tau^+$	$LF$	< 2.8 $\times 10^{-5}$	CL=90%
$\Gamma_{542}$	$K^+ \mu^\pm \tau^\mp$	$LF$	< 4.8 $\times 10^{-5}$	CL=90%
$\Gamma_{543}$	$K^*(892)^+ e^+ \mu^-$	$LF$	< 1.3 $\times 10^{-6}$	CL=90%
$\Gamma_{544}$	$K^*(892)^+ e^- \mu^+$	$LF$	< 9.9 $\times 10^{-7}$	CL=90%
$\Gamma_{545}$	$K^*(892)^+ e^\pm \mu^\mp$	$LF$	< 1.4 $\times 10^{-6}$	CL=90%
$\Gamma_{546}$	$\pi^- e^+ e^+$	$L$	< 2.3 $\times 10^{-8}$	CL=90%
$\Gamma_{547}$	$\pi^- \mu^+ \mu^+$	$L$	< 4.0 $\times 10^{-9}$	CL=95%
$\Gamma_{548}$	$\pi^- e^+ \mu^+$	$L$	< 1.5 $\times 10^{-7}$	CL=90%
$\Gamma_{549}$	$\rho^- e^+ e^+$	$L$	< 1.7 $\times 10^{-7}$	CL=90%
$\Gamma_{550}$	$\rho^- \mu^+ \mu^+$	$L$	< 4.2 $\times 10^{-7}$	CL=90%
$\Gamma_{551}$	$\rho^- e^+ \mu^+$	$L$	< 4.7 $\times 10^{-7}$	CL=90%
$\Gamma_{552}$	$K^- e^+ e^+$	$L$	< 3.0 $\times 10^{-8}$	CL=90%
$\Gamma_{553}$	$K^- \mu^+ \mu^+$	$L$	< 4.1 $\times 10^{-8}$	CL=90%
$\Gamma_{554}$	$K^- e^+ \mu^+$	$L$	< 1.6 $\times 10^{-7}$	CL=90%
$\Gamma_{555}$	$K^*(892)^- e^+ e^+$	$L$	< 4.0 $\times 10^{-7}$	CL=90%
$\Gamma_{556}$	$K^*(892)^- \mu^+ \mu^+$	$L$	< 5.9 $\times 10^{-7}$	CL=90%
$\Gamma_{557}$	$K^*(892)^- e^+ \mu^+$	$L$	< 3.0 $\times 10^{-7}$	CL=90%
$\Gamma_{558}$	$D^- e^+ e^+$	$L$	< 2.6 $\times 10^{-6}$	CL=90%
$\Gamma_{559}$	$D^- e^+ \mu^+$	$L$	< 1.8 $\times 10^{-6}$	CL=90%
$\Gamma_{560}$	$D^- \mu^+ \mu^+$	$L$	< 6.9 $\times 10^{-7}$	CL=95%
$\Gamma_{561}$	$D^{*-} \mu^+ \mu^+$	$L$	< 2.4 $\times 10^{-6}$	CL=95%
$\Gamma_{562}$	$D_c^- \mu^+ \mu^+$	$L$	< 5.8 $\times 10^{-7}$	CL=95%
$\Gamma_{563}$	$\bar{D}^0 \pi^- \mu^+ \mu^+$	$L$	< 1.5 $\times 10^{-6}$	CL=95%

# Meson Particle Listings

$B^\pm$

$\Gamma_{564}$	$\Lambda^0 \mu^+$	$L, B$	$< 6$	$\times 10^{-8}$	CL=90%
$\Gamma_{565}$	$\Lambda^0 e^+$	$L, B$	$< 3.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{566}$	$\bar{\Lambda}^0 \mu^+$	$L, B$	$< 6$	$\times 10^{-8}$	CL=90%
$\Gamma_{567}$	$\bar{\Lambda}^0 e^+$	$L, B$	$< 8$	$\times 10^{-8}$	CL=90%

- [a] An  $\ell$  indicates an  $e$  or a  $\mu$  mode, not a sum over these modes.  
 [b] An  $CP(\pm 1)$  indicates the  $CP=+1$  and  $CP=-1$  eigenstates of the  $D^0$ - $\bar{D}^0$  system.  
 [c]  $D$  denotes  $D^0$  or  $\bar{D}^0$ .  
 [d]  $D_{CP+}^{*0}$  decays into  $D^0 \pi^0$  with the  $D^0$  reconstructed in  $CP$ -even eigenstates  $K^+ K^-$  and  $\pi^+ \pi^-$ .  
 [e]  $\bar{D}^{**}$  represents an excited state with mass  $2.2 < M < 2.8$  GeV/ $c^2$ .  
 [f]  $X(3872)^+$  is a hypothetical charged partner of the  $X(3872)$ .  
 [g]  $\Theta(1710)^{++}$  is a possible narrow pentaquark state and  $G(2220)$  is a possible glueball resonance.  
 [h]  $(\bar{\Lambda}_c^- p)_s$  denotes a low-mass enhancement near 3.35 GeV/ $c^2$ .

## CONSTRAINED FIT INFORMATION

An overall fit to 3 branching ratios uses 6 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 3.7$  for 4 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_{365}$	$\frac{10}{x_{305}}$
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## CONSTRAINED FIT INFORMATION

An overall fit to 18 branching ratios uses 53 measurements and one constraint to determine 12 parameters. The overall fit has a  $\chi^2 = 49.2$  for 42 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_7$	33									
$x_{48}$	0	0								
$x_{98}$	0	0	8							
$x_{135}$	0	0	1	13						
$x_{255}$	0	0	0	0	0					
$x_{260}$	0	0	0	0	0	0				
$x_{273}$	0	0	0	0	0	28	0			
$x_{283}$	0	0	0	0	0	58	0	16		
$x_{516}$	0	0	0	0	0	14	0	4	8	
$x_{521}$	0	0	0	0	0	0	5	0	0	0
	$x_6$	$x_7$	$x_{48}$	$x_{98}$	$x_{135}$	$x_{255}$	$x_{260}$	$x_{273}$	$x_{283}$	$x_{516}$

## $B^+$ BRANCHING RATIOS

$\Gamma(\ell^+ \nu_\ell \text{ anything}) / \Gamma_{\text{total}}$   $\Gamma_1 / \Gamma$

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>10.99 ± 0.28 OUR EVALUATION</b>			
<b>10.76 ± 0.32 OUR AVERAGE</b>			Error includes scale factor of 1.1.
11.17 ± 0.25 ± 0.28	<sup>1</sup> URQUIJO	07	BELL $e^+ e^- \rightarrow \tau(4S)$
10.28 ± 0.26 ± 0.39	<sup>2</sup> AUBERT, B	06Y	BABR $e^+ e^- \rightarrow \tau(4S)$
10.25 ± 0.57 ± 0.65	<sup>3</sup> ARTUSO	97	CLE2 $e^+ e^- \rightarrow \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11.15 ± 0.26 ± 0.41	<sup>4</sup> OKABE	05	BELL Repl. by URQUIJO 07
10.1 ± 1.8 ± 1.5	ATHANAS	94	CLE2 Sup. by ARTUSO 97

<sup>1</sup> URQUIJO 07 report a measurement of  $(10.34 \pm 0.23 \pm 0.25)\%$  for the partial branching fraction of  $B^+ \rightarrow e^+ \nu_e X_C$  decay with electron energy above 0.6 GeV. We converted the result to  $B^+ \rightarrow e^+ \nu_e X$  branching fraction.

<sup>2</sup> The measurements are obtained for charged and neutral  $B$  mesons partial rates of semi-leptonic decay to electrons with momentum above 0.6 GeV/ $c$  in the  $B$  rest frame. The best precision on the ratio is achieved for a momentum threshold of 1.0 GeV:  $B(B^+ \rightarrow e^+ \nu_e X) / B(B^0 \rightarrow e^+ \nu_e X) = 1.074 \pm 0.041 \pm 0.026$ .

<sup>3</sup> ARTUSO 97 uses partial reconstruction of  $B \rightarrow D^* \ell \nu_\ell$  and inclusive semileptonic branching ratio from BARISH 96b ( $0.1049 \pm 0.0017 \pm 0.0043$ ).

<sup>4</sup> The measurements are obtained for charged and neutral  $B$  mesons partial rates of semi-leptonic decay to electrons with momentum above 0.6 GeV/ $c$  in the  $B$  rest frame, and their ratio of  $B(B^+ \rightarrow e^+ \nu_e X) / B(B^0 \rightarrow e^+ \nu_e X) = 1.08 \pm 0.05 \pm 0.02$ .

$\Gamma(e^+ \nu_e X_C) / \Gamma_{\text{total}}$   $\Gamma_2 / \Gamma$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>10.79 ± 0.25 ± 0.27</b>	<sup>1</sup> URQUIJO	07	BELL $e^+ e^- \rightarrow \tau(4S)$

<sup>1</sup> Measure the independent  $B^+$  and  $B^0$  partial branching fractions with electron threshold energies of 0.4 GeV.

$\Gamma(\bar{D}^0 \ell^+ \nu_\ell) / \Gamma_{\text{total}}$   $\Gamma_4 / \Gamma$

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.  $\ell = e$  or  $\mu$ , not sum over  $e$  and  $\mu$  modes.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0227 ± 0.0011 OUR EVALUATION</b>			
<b>0.0229 ± 0.0008 OUR AVERAGE</b>			
0.0229 ± 0.0008 ± 0.0009	<sup>1</sup> AUBERT	10	BABR $e^+ e^- \rightarrow \tau(4S)$
0.0234 ± 0.0003 ± 0.0013	AUBERT	09A	BABR $e^+ e^- \rightarrow \tau(4S)$
0.0221 ± 0.0013 ± 0.0019	<sup>2</sup> BARTELT	99	CLE2 $e^+ e^- \rightarrow \tau(4S)$
0.016 ± 0.006 ± 0.003	<sup>3</sup> FULTON	91	CLEO $e^+ e^- \rightarrow \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.0233 ± 0.0009 ± 0.0009	<sup>1</sup> AUBERT	08Q	BABR Repl. by AUBERT 09A
0.0194 ± 0.0015 ± 0.0034	<sup>4</sup> ATHANAS	97	CLE2 Repl. by BARTELT 99

<sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\tau(4S)$ .

<sup>3</sup> FULTON 91 assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at the  $\tau(4S)$ .

<sup>4</sup> ATHANAS 97 uses missing energy and missing momentum to reconstruct neutrino.

$\Gamma(\bar{D}^0 \ell^+ \nu_\ell) / \Gamma(\ell^+ \nu_\ell \text{ anything})$   $\Gamma_4 / \Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.255 ± 0.009 ± 0.009</b>	<sup>1</sup> AUBERT	10	BABR $e^+ e^- \rightarrow \tau(4S)$

<sup>1</sup> Uses a fully reconstructed  $B$  meson on the recoil side.

$\Gamma(\bar{D}^0 \ell^+ \nu_\ell) / \Gamma(D \ell^+ \nu_\ell \text{ anything})$   $\Gamma_4 / \Gamma_3$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.227 ± 0.014 ± 0.016</b>	<sup>1</sup> AUBERT	07AN	BABR $e^+ e^- \rightarrow \tau(4S)$

<sup>1</sup> Uses a fully reconstructed  $B$  meson on the recoil side.

$\Gamma(\bar{D}^0 \tau^+ \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_5 / \Gamma$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.77 ± 0.22 ± 0.12</b>	<sup>1</sup> BOZEK	10	BELL $e^+ e^- \rightarrow \tau(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.67 ± 0.37 ± 0.13 <sup>2</sup> AUBERT 08N BABR Repl. by AUBERT 09s

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\tau(4S)$ .

<sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

$\Gamma(\bar{D}^0 \tau^+ \nu_\tau) / \Gamma(\bar{D}^0 \ell^+ \nu_\ell)$   $\Gamma_5 / \Gamma_4$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.429 ± 0.082 ± 0.052</b>	<sup>1,2</sup> LEES	12D	BABR $e^+ e^- \rightarrow \tau(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.314 ± 0.170 ± 0.049 <sup>1</sup> AUBERT 09s BABR Repl. by LEES 12D

<sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

<sup>2</sup> Uses  $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$  and  $\tau^+ \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$  and  $e^+$  or  $\mu^+$  as  $\ell^+$ .

$\Gamma(\bar{D}^{*}(2007)^0 \ell^+ \nu_\ell) / \Gamma_{\text{total}}$   $\Gamma_6 / \Gamma$

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.  $\ell = e$  or  $\mu$ , not sum over  $e$  and  $\mu$  modes.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0569 ± 0.0019 OUR EVALUATION</b>				
<b>0.0560 ± 0.0026 OUR FIT</b>				Error includes scale factor of 1.5.
<b>0.0558 ± 0.0026 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
0.0540 ± 0.0002 ± 0.0021		AUBERT	09A	BABR $e^+ e^- \rightarrow \tau(4S)$
0.0556 ± 0.0008 ± 0.0041	<sup>1</sup> AUBERT	08AT	BABR	$e^+ e^- \rightarrow \tau(4S)$
0.0650 ± 0.0020 ± 0.0043	<sup>2</sup> ADAM	03	CLE2	$e^+ e^- \rightarrow \tau(4S)$
0.066 ± 0.016 ± 0.015	<sup>3</sup> ALBRECHT	92C	ARG	$e^+ e^- \rightarrow \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0583 ± 0.0015 ± 0.0030	<sup>4</sup> AUBERT	08Q	BABR	Repl. by AUBERT 09A
0.0650 ± 0.0020 ± 0.0043	<sup>5</sup> BRIERE	02	CLE2	$e^+ e^- \rightarrow \tau(4S)$
0.0513 ± 0.0054 ± 0.0064	<sup>6</sup> BARISH	95	CLE2	Repl. by ADAM 03
seen	<sup>7</sup> SANGHERA	93	CLE2	$e^+ e^- \rightarrow \tau(4S)$
0.041 ± 0.008 ± 0.008	<sup>8</sup> FULTON	91	CLEO	$e^+ e^- \rightarrow \tau(4S)$
0.070 ± 0.018 ± 0.014	<sup>9</sup> ANTREASNYAN	90B	CBAL	$e^+ e^- \rightarrow \tau(4S)$

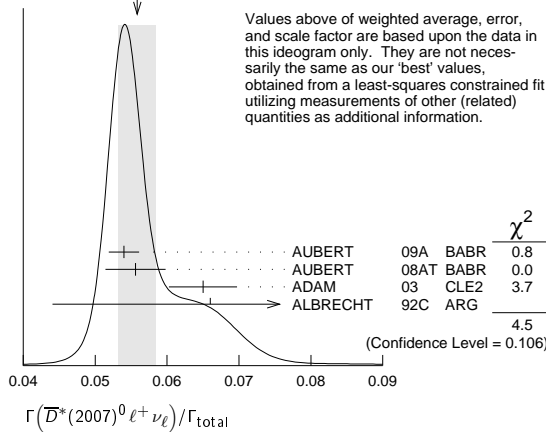
See key on page 601

## Meson Particle Listings

 $B^\pm$ 

- <sup>1</sup> Measured using the dependence of  $B^- \rightarrow D^{*0} e^- \bar{\nu}_e$  decay differential rate and the form factor description by CAPRINI 98.
- <sup>2</sup> Simultaneous measurements of both  $B^0 \rightarrow D^*(2010)^- \ell \nu$  and  $B^+ \rightarrow \bar{D}(2007)^0 \ell \nu$ .
- <sup>3</sup> ALBRECHT 92C reports  $0.058 \pm 0.014 \pm 0.013$ . We rescale using the method described in STONE 94 but with the updated PDG 94  $B(D^0 \rightarrow K^- \pi^+)$ . Assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at the  $\Upsilon(4S)$ .
- <sup>4</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.
- <sup>5</sup> The results are based on the same analysis and data sample reported in ADAM 03.
- <sup>6</sup> BARISH 95 use  $B(D^0 \rightarrow K^- \pi^+) = (3.91 \pm 0.08 \pm 0.17)\%$  and  $B(D^{*0} \rightarrow D^0 \pi^0) = (63.6 \pm 2.3 \pm 3.3)\%$ .
- <sup>7</sup> Combining  $\bar{D}^{*0} \ell^+ \nu_\ell$  and  $\bar{D}^{*-} \ell^+ \nu_\ell$  SANGHERA 93 test  $V-A$  structure and fit the decay angular distributions to obtain  $A_{FB} = 3/4 * (\Gamma^- - \Gamma^+)/\Gamma = 0.14 \pm 0.06 \pm 0.03$ . Assuming a value of  $V_{cb}$ , they measure  $V, A_1$ , and  $A_2$ , the three form factors for the  $D^* \ell \nu_\ell$  decay, where results are slightly dependent on model assumptions.
- <sup>8</sup> Assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at the  $\Upsilon(4S)$ . Uncorrected for  $D$  and  $D^*$  branching ratio assumptions.
- <sup>9</sup> ANTREASYAN 90B is average over  $B$  and  $\bar{D}^*(2010)$  charge states.

WEIGHTED AVERAGE  
 $0.0558 \pm 0.0026$  (Error scaled by 1.5)



$\Gamma(\bar{D}^*(2007)^0 \ell^+ \nu_\ell) / \Gamma(D \ell^+ \nu_\ell \text{ anything})$				$\Gamma_6 / \Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.582 \pm 0.018 \pm 0.030</math></b>	<sup>1</sup> AUBERT	07AN BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses a fully reconstructed  $B$  meson on the recoil side.

$\Gamma(\bar{D}^*(2007)^0 \tau^+ \nu_\tau) / \Gamma_{\text{total}}$				$\Gamma_7 / \Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.88 \pm 0.20</math> OUR FIT</b>				
<b><math>2.12^{+0.28}_{-0.27} \pm 0.29</math></b>	<sup>1</sup> BOZEK	10 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.25 \pm 0.48 \pm 0.28$	<sup>2</sup> AUBERT	08N BABR	Repl. by AUBERT 09S	

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

$\Gamma(\bar{D}^*(2007)^0 \tau^+ \nu_\tau) / \Gamma(\bar{D}^*(2007)^0 \ell^+ \nu_\ell)$				$\Gamma_7 / \Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.335 \pm 0.034</math> OUR FIT</b>				
<b><math>0.322 \pm 0.032 \pm 0.022</math></b>	<sup>1,2</sup> LEES	12D BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.346 \pm 0.073 \pm 0.034$	<sup>1</sup> AUBERT	09S BABR	Repl. by LEES 12D	

- <sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.
- <sup>2</sup> Uses  $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$  and  $\tau^- \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$  and  $e^+$  or  $\mu^+$  as  $\ell^+$ .

$\Gamma(D^- \pi^+ \ell^+ \nu_\ell) / \Gamma_{\text{total}}$				$\Gamma_8 / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>4.2 \pm 0.5</math> OUR AVERAGE</b>				
$4.2 \pm 0.6 \pm 0.3$	<sup>1</sup> AUBERT	08Q BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$4.2 \pm 0.6 \pm 0.2$	<sup>1,2</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$5.5 \pm 0.9 \pm 0.3$	<sup>3</sup> LIVENTSEV	05 BELL	Repl. by LIVENTSEV 08	

- <sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.
- <sup>2</sup> LIVENTSEV 08 reports  $(4.0 \pm 0.4 \pm 0.6) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow D^- \pi^+ \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell)]$  assuming  $B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell) = (2.15 \pm 0.22) \times 10^{-2}$ , which we rescale to our best value  $B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell) = (2.27 \pm 0.11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>3</sup> LIVENTSEV 05 reports  $[\Gamma(B^+ \rightarrow D^- \pi^+ \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \ell^+ \nu_\ell)] = 0.25 \pm 0.03 \pm 0.03$  which we multiply by our best value  $B(B^0 \rightarrow D^- \ell^+ \nu_\ell) = (2.19 \pm 0.12) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{D}_2^0(2420)^0 \ell^+ \nu_\ell, \bar{D}_2^0 \rightarrow D^- \pi^+) / \Gamma_{\text{total}}$				$\Gamma_9 / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>2.5 \pm 0.5</math> OUR AVERAGE</b>				
$2.6 \pm 0.5 \pm 0.4$	<sup>1</sup> AUBERT	08BL BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$2.4 \pm 0.4 \pm 0.6$	<sup>1</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

$\Gamma(\bar{D}_2^*(2460)^0 \ell^+ \nu_\ell, \bar{D}_2^{*0} \rightarrow D^- \pi^+) / \Gamma_{\text{total}}$				$\Gamma_{10} / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.53 \pm 0.16</math> OUR AVERAGE</b>				
$1.42 \pm 0.15 \pm 0.15$	<sup>1</sup> AUBERT	09Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$1.5 \pm 0.2 \pm 0.2$	<sup>2</sup> AUBERT	08BL BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$2.2 \pm 0.3 \pm 0.4$	<sup>2</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

- <sup>1</sup> Uses a simultaneous fit of all  $B$  semileptonic decays without full reconstruction of events. AUBERT 09Y reports  $B(B^+ \rightarrow \bar{D}_2^*(2460)^0 \ell^+ \nu_\ell) \cdot B(\bar{D}_2^*(2460)^0 \rightarrow D^{*-} \pi^+) = (2.29 \pm 0.23 \pm 0.21) \times 10^{-3}$  and the authors have provided us the individual measurement.
- <sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

$\Gamma(D^{*-} n \pi \ell^+ \nu_\ell (n \geq 1)) / \Gamma(D \ell^+ \nu_\ell \text{ anything})$				$\Gamma_{11} / \Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.191 \pm 0.013 \pm 0.019</math></b>	<sup>1</sup> AUBERT	07AN BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses a fully reconstructed  $B$  meson on the recoil side.

$\Gamma(D^{*-} \pi^+ \ell^+ \nu_\ell) / \Gamma_{\text{total}}$				$\Gamma_{12} / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>6.1 \pm 0.6</math> OUR AVERAGE</b>				
$5.9 \pm 0.5 \pm 0.4$	<sup>1</sup> AUBERT	08Q BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$6.8 \pm 1.1 \pm 0.3$	<sup>1,2</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$5.9 \pm 1.4 \pm 0.1$	<sup>3,4</sup> LIVENTSEV	05 BELL	Repl. by LIVENTSEV 08	

- <sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.
- <sup>2</sup> LIVENTSEV 08 reports  $(6.4 \pm 0.8 \pm 0.9) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow D^{*-} \pi^+ \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell)]$  assuming  $B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell) = (2.15 \pm 0.22) \times 10^{-2}$ , which we rescale to our best value  $B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell) = (2.27 \pm 0.11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>3</sup> Excludes  $D^{*+}$  contribution to  $D \pi$  modes.
- <sup>4</sup> LIVENTSEV 05 reports  $[\Gamma(B^+ \rightarrow D^{*-} \pi^+ \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow D^*(2010)^- \ell^+ \nu_\ell)] = 0.12 \pm 0.02 \pm 0.02$  which we multiply by our best value  $B(B^0 \rightarrow D^*(2010)^- \ell^+ \nu_\ell) = (4.93 \pm 0.11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{D}_1(2420)^0 \ell^+ \nu_\ell, \bar{D}_1^0 \rightarrow D^{*-} \pi^+) / \Gamma_{\text{total}}$				$\Gamma_{13} / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>3.03 \pm 0.20</math> OUR AVERAGE</b>				
$2.97 \pm 0.17 \pm 0.17$	<sup>1</sup> AUBERT	09Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$2.9 \pm 0.3 \pm 0.3$	<sup>2</sup> AUBERT	08BL BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$4.2 \pm 0.7 \pm 0.7$	<sup>2</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
$3.73 \pm 0.85 \pm 0.57$	<sup>3</sup> ANASTASSOV	98 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	

- <sup>1</sup> Uses a simultaneous measurement of all  $B$  semileptonic decays without full reconstruction of events.
- <sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.
- <sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}_1^*(2430)^0 \ell^+ \nu_\ell, \bar{D}_1^{*0} \rightarrow D^{*-} \pi^+) / \Gamma_{\text{total}}$				$\Gamma_{14} / \Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.7 \pm 0.4 \pm 0.5</math></b>		<sup>1</sup> AUBERT	08BL BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.7$	90	<sup>1</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

$\Gamma(\bar{D}_2^*(2460)^0 \ell^+ \nu_\ell, \bar{D}_2^{*0} \rightarrow D^{*-} \pi^+) / \Gamma_{\text{total}}$				$\Gamma_{15} / \Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.01 \pm 0.24</math> OUR AVERAGE</b>				Error includes scale factor of 2.0.
$0.87 \pm 0.11 \pm 0.07$		<sup>1</sup> AUBERT	09Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.5 \pm 0.2 \pm 0.2$		<sup>2</sup> AUBERT	08BL BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.8 \pm 0.6 \pm 0.3$		<sup>2</sup> LIVENTSEV	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.6$	90	<sup>3</sup> ANASTASSOV	98 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Uses a simultaneous fit of all  $B$  semileptonic decays without full reconstruction of events. AUBERT 09Y reports  $B(B^+ \rightarrow \bar{D}_2^*(2460)^0 \ell^+ \nu_\ell) \cdot B(\bar{D}_2^*(2460)^0 \rightarrow D^{*-} \pi^+) = (2.29 \pm 0.23 \pm 0.21) \times 10^{-3}$  and the authors have provided us the individual measurement.
- <sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.
- <sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0 \pi^+ \pi^- \ell^+ \nu_\ell) / \Gamma(\bar{D}^0 \ell^+ \nu_\ell)$				$\Gamma_{16} / \Gamma_4$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>7.1 \pm 1.3 \pm 0.8</math></b>	<sup>1</sup> LEES	16 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Measurement used electrons and muons as leptons.

# Meson Particle Listings

$B^\pm$

$\Gamma(\bar{D}^{*0}\pi^+\pi^-\ell^+\nu_\ell)/\Gamma(\bar{D}^{*0}\ell^+\nu_\ell)$	$\Gamma_{17}/\Gamma_6$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.4 \pm 0.7 \pm 0.4</math></b>	<sup>1</sup> LEES	16	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Measurement used electrons and muons as leptons.				

$\Gamma(D_s^{(*)-}K^+\ell^+\nu_\ell)/\Gamma_{\text{total}}$	$\Gamma_{18}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>6.1 \pm 1.0</math> OUR AVERAGE</b>				
$5.9 \pm 1.2 \pm 1.5$	<sup>1</sup> STYPULA	12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$6.13^{+1.04}_{-1.03} \pm 0.67$	<sup>1</sup> DEL-AMO-SA..11L	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_s^-K^+\ell^+\nu_\ell)/\Gamma_{\text{total}}$	$\Gamma_{19}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>3.0 \pm 0.9 \pm 1.1</math></b>	<sup>1</sup> STYPULA	12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_s^{*-}K^+\ell^+\nu_\ell)/\Gamma_{\text{total}}$	$\Gamma_{20}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>2.9 \pm 1.6 \pm 1.1</math></b>	<sup>1,2</sup> STYPULA	12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>2</sup> STYPULA 12 provides also an upper limit of $0.56 \times 10^{-3}$ at 90% CL for the same data. Also measures branching fraction of the combined modes of $D_s^-K^+\ell^+\nu_\ell$ and $D_s^{*-}K^+\ell^+\nu_\ell$ as $B(B^+ \rightarrow D_s^{(*)-}K^+\ell^+\nu_\ell) = (5.9 \pm 1.2 \pm 1.5) \times 10^{-4}$ .				

$\Gamma(\pi^0\ell^+\nu_\ell)/\Gamma_{\text{total}}$	$\Gamma_{21}/\Gamma$			
“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <a href="http://www.slac.stanford.edu/xorg/hfag/">http://www.slac.stanford.edu/xorg/hfag/</a> . The averaging/rescaling procedure takes into account correlations between the measurements.				
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.780 \pm 0.027</math> OUR EVALUATION</b>				
<b><math>0.748 \pm 0.029</math> OUR AVERAGE</b>				
$0.80 \pm 0.08 \pm 0.04$	<sup>1</sup> SIBIDANOV	13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.77 \pm 0.04 \pm 0.03$	<sup>2</sup> LEES	12AA	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.705 \pm 0.025 \pm 0.035$	<sup>3</sup> DEL-AMO-SA..11c	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
$0.82 \pm 0.09 \pm 0.05$	<sup>4</sup> AUBERT	08AV	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.77 \pm 0.14 \pm 0.08$	<sup>4</sup> HOKUUE	07	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.74 \pm 0.05 \pm 0.10$	<sup>5</sup> AUBERT,B	05o	BABR	Repl. by DEL-AMO-SANCHEZ 11c
<sup>1</sup> The signal events are tagged by a second $B$ meson reconstructed in the fully hadronic decays. <sup>2</sup> Uses loose neutrino reconstruction technique. Assumes $B(Y(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$ and $B(Y(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ . <sup>3</sup> Using the isospin symmetry relation, $B^+$ and $B^0$ branching fractions are combined. <sup>4</sup> The signal events are tagged by a second $B$ meson reconstructed in the semileptonic mode $B \rightarrow D^{(*)}\ell\nu_\ell$ . <sup>5</sup> $B^+$ and $B^0$ decays combined assuming isospin symmetry. Systematic errors include both experimental and form-factor uncertainties.				

$\Gamma(\pi^0e^+\nu_e)/\Gamma_{\text{total}}$	$\Gamma_{22}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.9 \pm 0.2 \pm 0.2$		<sup>1</sup> ALEXANDER	96T	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<22	90	ANTREASYAN	90B	CBAL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Derived based in the reported $B^0$ result by assuming isospin symmetry: $\Gamma(B^0 \rightarrow \pi^-\ell^+\nu) = 2\Gamma(B^+ \rightarrow \pi^0\ell^+\nu)$ .				

$\Gamma(\eta\ell^+\nu_\ell)/\Gamma_{\text{total}}$	$\Gamma_{23}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.38 \pm 0.06</math> OUR AVERAGE</b>				
$0.38 \pm 0.05 \pm 0.05$		<sup>1</sup> LEES	12AA	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.31 \pm 0.06 \pm 0.08$		<sup>1</sup> AUBERT	09Q	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.64 \pm 0.20 \pm 0.03$		<sup>2</sup> AUBERT	08AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.36 \pm 0.05 \pm 0.04$		<sup>1</sup> DEL-AMO-SA..11F	BABR	Repl. by LEES 12AA
<1.01	90	<sup>3</sup> ADAM	07	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$0.84 \pm 0.31 \pm 0.18$		<sup>4</sup> ATHAR	03	CLE2 Repl. by ADAM 07
<sup>1</sup> Uses loose neutrino reconstruction technique. Assumes $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ . <sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>3</sup> The $B^0$ and $B^+$ results are combined assuming the isospin, $B$ lifetimes, and relative charged/neutral $B$ production at the $\Upsilon(4S)$ . <sup>4</sup> ATHAR 03 reports systematic errors $0.16 \pm 0.09$ , which are experimental systematic and systematic due to model dependence. We combine these in quadrature.				

$\Gamma(\eta'\ell^+\nu_\ell)/\Gamma_{\text{total}}$	$\Gamma_{24}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.23 \pm 0.08</math> OUR AVERAGE</b>				
$0.24 \pm 0.08 \pm 0.03$	<sup>1</sup> LEES	12AA	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.04 \pm 0.22^{+0.05}_{-0.02}$	<sup>2</sup> AUBERT	08AV	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$2.66 \pm 0.80 \pm 0.56$	<sup>3</sup> ADAM	07	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.24 \pm 0.08 \pm 0.03$	<sup>1</sup> DEL-AMO-SA..11F	BABR	Repl. by LEES 12AA	
<sup>1</sup> Uses loose neutrino reconstruction technique. Assumes $B(Y(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$ and $B(Y(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ . <sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>3</sup> The $B^0$ and $B^+$ results are combined assuming the isospin, $B$ lifetimes, and relative charged/neutral $B$ production at the $\Upsilon(4S)$ . Corresponds to 90% CL interval $(1.20\text{--}4.46) \times 10^{-4}$ .				

$\Gamma(\omega\ell^+\nu_\ell)/\Gamma_{\text{total}}$ $\ell = e \text{ or } \mu$ , not sum over $e$ and $\mu$ modes.	$\Gamma_{25}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.19 ± 0.09 OUR AVERAGE</b>				
1.21 ± 0.14 ± 0.08	<sup>1,2</sup>	LEES	13A	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.35 ± 0.21 ± 0.11	<sup>3</sup>	LEES	13T	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.07 ± 0.16 ± 0.07	<sup>4</sup>	SIBIDANOV	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
1.19 ± 0.16 ± 0.09	<sup>2,5</sup>	LEES	12AA	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.3 ± 0.4 ± 0.4	<sup>6</sup>	SCHWANDA	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.14 ± 0.16 ± 0.08	<sup>2</sup>	AUBERT	09Q	BABR Repl. by LEES 13A
<2.1	90	<sup>7</sup> BEAN	93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> LEES 13A reports $(1.21 \pm 0.14 \pm 0.08) \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow \omega\ell^+\nu_\ell)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow \pi^+\pi^-\pi^0)]$ assuming $B(\omega(782) \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7) \times 10^{-2}$ .				
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .				
<sup>3</sup> Uses semileptonic tagging. Assumes $B(\omega \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7)\%$ and that the production ratio of $B^+B^-$ to $B^0\bar{B}^0$ from $\Upsilon(4S)$ is $1.056 \pm 0.028$ . The partial branching fractions in three bins of $q^2$ are also reported.				
<sup>4</sup> The signal events are tagged by a second $B$ meson reconstructed in the fully hadronic decays.				
<sup>5</sup> Uses loose neutrino reconstruction technique.				
<sup>6</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>7</sup> BEAN 93B limit set using ISGW Model. Using isospin and the quark model to combine $\Gamma(\rho^0\ell^+\nu_\ell)$ and $\Gamma(\rho^-\ell^+\nu_\ell)$ with this result, they obtain a limit $<(1.6\text{--}2.7) \times 10^{-4}$ at 90% CL for $B^+ \rightarrow \omega\ell^+\nu_\ell$ . The range corresponds to the ISGW, WSB, and KS models.				
An upper limit on $ V_{ub}/V_{cb}  < 0.8\text{--}0.13$ at 90% CL is derived as well.				

$\Gamma(\omega\mu^+\nu_\mu)/\Gamma_{\text{total}}$	$\Gamma_{26}/\Gamma$		
VALUE	DOCUMENT ID	TECN	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
seen	<sup>1</sup> ALBRECHT	91c	ARG
<sup>1</sup> In ALBRECHT 91c, one event is fully reconstructed providing evidence for the $b \rightarrow u$ transition.			

$\Gamma(\rho^0 \ell^+ \nu_\ell) / \Gamma_{\text{total}}$   
 $\ell = e \text{ or } \mu$ , not sum over  $e$  and  $\mu$  modes.

$\Gamma_{27}/\Gamma$

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.58 \pm 0.11</math> OUR EVALUATION</b>				
<b><math>1.42 \pm 0.23</math> OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.			
$1.83 \pm 0.10 \pm 0.10$		<sup>1</sup> SIBIDANOV	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.94 \pm 0.08 \pm 0.14$		<sup>2</sup> DEL-AMO-SA..11c	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.33 \pm 0.23 \pm 0.18$		<sup>3</sup> HOKUUE	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$1.34 \pm 0.15^{+0.28}_{-0.32}$		<sup>4</sup> BEHRENS	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.16 \pm 0.11 \pm 0.30$		<sup>2</sup> AUBERT,B	05o	BABR Repl. by DEL-AMO-SANCHEZ 11c
$1.40 \pm 0.21^{+0.32}_{-0.33}$		<sup>4</sup> BEHRENS	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$1.2 \pm 0.2^{+0.3}_{-0.4}$		<sup>4</sup> ALEXANDER	96T	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<2.1	90	<sup>5</sup> BEAN	93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> The signal events are tagged by a second  $B$  meson reconstructed in the fully hadronic decays.

<sup>2</sup>  $B^+$  and  $B^0$  decays combined assuming isospin symmetry. Systematic errors include both experimental and form-factor uncertainties.

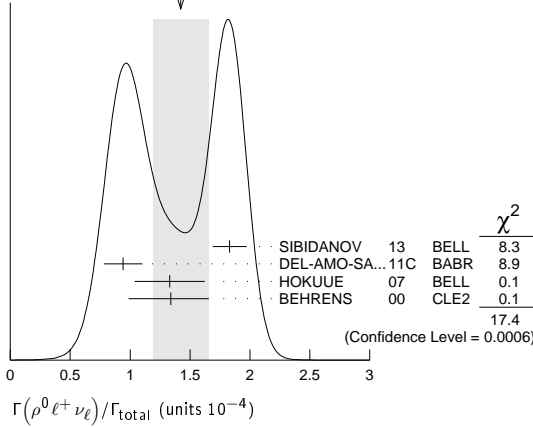
<sup>3</sup> The signal events are tagged by a second  $B$  meson reconstructed in the semileptonic mode  $B \rightarrow D^{(*)} \ell \nu_\ell$ .

<sup>4</sup> Derived based in the reported  $B^0$  result by assuming isospin symmetry:  $\Gamma(B^0 \rightarrow \rho^- \ell^+ \nu) = 2\Gamma(B^+ \rightarrow \rho^0 \ell^+ \nu) \approx 2\Gamma(B^+ \rightarrow \omega \ell^+ \nu)$ .

<sup>5</sup> BEAN 93B limit set using ISGW Model. Using isospin and the quark model to combine  $\Gamma(\omega^0 \ell^+ \nu_\ell)$  and  $\Gamma(\rho^- \ell^+ \nu_\ell)$  with this result, they obtain a limit  $<(1.6\text{--}2.7) \times 10^{-4}$

at 90% CL for  $B^+ \rightarrow \rho^0 \ell^+ \nu_\ell$ . The range corresponds to the ISGW, WSB, and KS models. An upper limit on  $|V_{ub}/V_{cb}| < 0.8\text{--}0.13$  at 90% CL is derived as well.

WEIGHTED AVERAGE  
1.42±0.23 (Error scaled by 2.4)



$\Gamma(\rho^0 \ell^+ \nu_\ell)/\Gamma_{\text{total}}$		$\Gamma_{28}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>5.8±2.1±0.9</b>	90	<sup>1</sup> TIEN	14 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\rho^0 \mu^+ \nu_\mu)/\Gamma_{\text{total}}$		$\Gamma_{29}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;8.5 × 10<sup>-6</sup></b>	90	<sup>1</sup> TIEN	14 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\rho^0 e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{30}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>8.2±3.7±0.6</b>	90	<sup>1</sup> TIEN	14 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
<5200	90	<sup>2</sup> ADAM	03B CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Based on phase-space model; if  $V\text{--}A$  model is used, the 90% CL upper limit becomes  $< 1.2 \times 10^{-3}$ .

$\Gamma(e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{31}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt; 0.98</b>	90	<sup>1</sup> SATOYAMA	07 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
< 3.5	90	<sup>2</sup> YOOK	15 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< 8	90	<sup>1</sup> AUBERT	10E BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 1.9	90	<sup>1</sup> AUBERT	09V BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 5.2	90	<sup>1</sup> AUBERT	08AD BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 15	90	ARTUSO	95 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Assumes  $B(\Upsilon(4S) \rightarrow B^+ B^-) = 0.513 \pm 0.006$ .

$\Gamma(\mu^+ \nu_\mu)/\Gamma_{\text{total}}$		$\Gamma_{32}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt; 1.0</b>	90	<sup>1</sup> AUBERT	09V BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
< 2.7	90	<sup>2</sup> YOOK	15 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< 11	90	<sup>1</sup> AUBERT	10E BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 5.6	90	<sup>1</sup> AUBERT	08AD BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 1.7	90	<sup>1</sup> SATOYAMA	07 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< 6.6	90	AUBERT	04O BABR Repl. by AUBERT 09V
< 21	90	ARTUSO	95 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Assumes  $B(\Upsilon(4S) \rightarrow B^+ B^-) = 0.513 \pm 0.006$ .

$\Gamma(\tau^+ \nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{33}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>1.09±0.24 OUR AVERAGE</b>		Error includes scale factor of 1.2.	
1.25±0.28±0.27	1.2	KRONENBIT...	15 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
0.72±0.27±0.11	3	HARA	13 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
1.83±0.53±0.24	2.4	LEES	13K BABR $e^+ e^- \rightarrow \Upsilon(4S)$
1.7 ± 0.8 ± 0.2	2.5	AUBERT	10E BABR $e^+ e^- \rightarrow \Upsilon(4S)$

See the note on "Decay Constants of Charged Pseudoscalar Mesons" in the  $D_s^+$  Listings.

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
1.54±0.38±0.29 -0.37 - 0.31		<sup>2,6</sup> HARA	10 BELL Repl. by KRONENBIT-TER 15
1.8 ± 0.9 ± 0.45		<sup>2,7</sup> AUBERT	08D BABR Repl. by LEES 13K
0.9 ± 0.6 ± 0.1		<sup>2,5</sup> AUBERT	07AL BABR Repl. by AUBERT 10E
< 2.6	90	<sup>2</sup> AUBERT	06K BABR $e^+ e^- \rightarrow \Upsilon(4S)$
1.79±0.56±0.46 -0.49 - 0.51		<sup>2,7</sup> IKADO	06 BELL Repl. by HARA 13
< 4.2	90	<sup>2</sup> AUBERT,B	05B BABR Repl. by AUBERT 06K
< 8.3	90	<sup>8</sup> BARATE	01E ALEP $e^+ e^- \rightarrow Z$
< 8.4	90	<sup>2</sup> BROWDER	01 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
< 5.7	90	<sup>9</sup> ACCIARRI	97F L3 $e^+ e^- \rightarrow Z$
< 104	90	<sup>10</sup> ALBRECHT	95D ARG $e^+ e^- \rightarrow \Upsilon(4S)$
< 22	90	ARTUSO	95D CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
< 18	90	<sup>11</sup> BUSKULIC	95 ALEP $e^+ e^- \rightarrow Z$

<sup>1</sup> Requires one reconstructed semileptonic  $B$  decay  $B^- \rightarrow D^{(*)0} \ell^- \bar{\nu}_\ell$  in the recoil.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> The authors combine their result with that from HARA 10 obtaining  $B(B^- \rightarrow \tau^- \bar{\nu}_\tau) = (0.96 \pm 0.26) \times 10^{-4}$  and deriving  $f_B |V_{ub}| = (7.4 \pm 0.8 \pm 0.5) \times 10^{-4}$  GeV.

<sup>4</sup> Requires a fully reconstructed hadronic  $B$ -decay in the recoil. Reports that this result combined with AUBERT 10E value gives  $B(B^- \rightarrow \tau^- \bar{\nu}_\tau) = (1.79 \pm 0.48) \times 10^{-4}$ .

<sup>5</sup> Requires one reconstructed semileptonic  $B$  decay  $B^- \rightarrow D^0 \ell^- \bar{\nu}_\ell X$  in the recoil.

<sup>6</sup> Requires one reconstructed semileptonic  $B$  decay  $B^- \rightarrow D^{(*)0} \ell^- \bar{\nu}_\ell X$  in the recoil.

<sup>7</sup> The analysis is based on a sample of events with one fully reconstructed tag  $B$  in a hadronic decay mode  $B^- \rightarrow D^{(*)0} X^-$ .

<sup>8</sup> The energy-flow and  $b$ -tagging algorithms were used.

<sup>9</sup> ACCIARRI 97F uses missing-energy technique and  $f(b \rightarrow B^-) = (38.2 \pm 2.5)\%$ .

<sup>10</sup> ALBRECHT 95D uses full reconstruction of one  $B$  decay as tag.

<sup>11</sup> BUSKULIC 95 uses same missing-energy technique as in  $\bar{B} \rightarrow \tau^+ \nu_\tau X$ , but analysis is restricted to endpoint region of missing-energy distribution.

$\Gamma(\ell^+ \nu_\ell \gamma)/\Gamma_{\text{total}}$		$\Gamma_{34}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt; 3.5 × 10<sup>-6</sup></b>	90	<sup>1</sup> HELLER	15 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
<15.6 × 10 <sup>-6</sup>	90	<sup>1</sup> AUBERT	09AT BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(e^+ \nu_e \gamma)/\Gamma_{\text{total}}$		$\Gamma_{35}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt; 6.1 × 10<sup>-6</sup></b>	90	<sup>1</sup> HELLER	15 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
< 17 × 10 <sup>-6</sup>	90	<sup>1</sup> AUBERT	09AT BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 200 × 10 <sup>-6</sup>	90	<sup>2</sup> BROWDER	97 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> BROWDER 97 uses the hermiticity of the CLEOII detector to reconstruct the neutrino energy and momentum.

$\Gamma(\mu^+ \nu_\mu \gamma)/\Gamma_{\text{total}}$		$\Gamma_{36}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt; 3.4 × 10<sup>-6</sup></b>	90	<sup>1</sup> HELLER	15 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
< 24 × 10 <sup>-6</sup>	90	<sup>1,2</sup> AUBERT	09AT BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 52 × 10 <sup>-6</sup>	90	<sup>3</sup> BROWDER	97 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Note that the value given by Aubert 2009 is 24 E-6 in the paper abstract, and 26 E-6 in the paper itself (Table I).

<sup>3</sup> BROWDER 97 uses the hermiticity of the CLEOII detector to reconstruct the neutrino energy and momentum.

$\Gamma(D^0 X)/\Gamma_{\text{total}}$		$\Gamma_{37}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>0.086±0.006±0.004</b>		<sup>1</sup> AUBERT	07N BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
0.098±0.009±0.006		<sup>1</sup> AUBERT, BE	04B BABR Repl. by AUBERT 07N

<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.

$\Gamma(D^0 X)/\Gamma_{\text{total}}$		$\Gamma_{38}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>0.786±0.016±0.034 -0.033</b>		<sup>1</sup> AUBERT	07N BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN COMMENT
0.793±0.025±0.045 -0.044		<sup>1</sup> AUBERT, BE	04B BABR Repl. by AUBERT 07N

<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.

# Meson Particle Listings

$B^\pm$

$\Gamma(D^0 X)/[\Gamma(D^0 X) + \Gamma(\overline{D}^0 X)]$	$\Gamma_{37}/(\Gamma_{37} + \Gamma_{38})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.098 \pm 0.007 \pm 0.001</math></b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.110 \pm 0.010 \pm 0.003$	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(D^+ X)/\Gamma_{\text{total}}$	$\Gamma_{39}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.025 \pm 0.005 \pm 0.002</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.038 \pm 0.009 \pm 0.005$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D^- X)/\Gamma_{\text{total}}$	$\Gamma_{40}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.099 \pm 0.008 \pm 0.009</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.098 \pm 0.012 \pm 0.014$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D^+ X)/[\Gamma(D^+ X) + \Gamma(D^- X)]$	$\Gamma_{39}/(\Gamma_{39} + \Gamma_{40})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.204 \pm 0.035 \pm 0.001</math></b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.278 \pm 0.052 \pm 0.009$	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(D_s^+ X)/\Gamma_{\text{total}}$	$\Gamma_{41}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.079 \pm 0.006 \pm \overset{+0.013}{-0.011}</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.143 \pm 0.016 \pm \overset{+0.051}{-0.034}$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D_s^- X)/\Gamma_{\text{total}}$	$\Gamma_{42}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.011 \pm 0.004 \pm \overset{+0.002}{-0.001}</math></b>		<sup>1</sup> AUBERT	07N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
<0.022	90	<sup>1</sup> AUBERT,BE	04B	BABR Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D_s^+ X)/[\Gamma(D_s^+ X) + \Gamma(D_s^- X)]$	$\Gamma_{41}/(\Gamma_{41} + \Gamma_{42})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.884 \pm 0.038 \pm 0.002</math></b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.966 \pm 0.039 \pm 0.012$	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(D_s^- X)/[\Gamma(D_s^+ X) + \Gamma(D_s^- X)]$	$\Gamma_{42}/(\Gamma_{41} + \Gamma_{42})$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.126</b>	90	AUBERT,BE	04B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(\Lambda_c^+ X)/\Gamma_{\text{total}}$	$\Gamma_{43}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.021 \pm 0.005 \pm \overset{+0.008}{-0.004}</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.029 \pm 0.008 \pm \overset{+0.011}{-0.007}$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(\overline{\Lambda}_c^- X)/\Gamma_{\text{total}}$	$\Gamma_{44}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.028 \pm 0.005 \pm \overset{+0.010}{-0.007}</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.035 \pm 0.008 \pm \overset{+0.013}{-0.009}$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(\Lambda_c^+ X)/[\Gamma(\Lambda_c^+ X) + \Gamma(\overline{\Lambda}_c^- X)]$	$\Gamma_{43}/(\Gamma_{43} + \Gamma_{44})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.427 \pm 0.071 \pm 0.001</math></b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.452 \pm 0.090 \pm 0.003$	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(\overline{\Sigma}^+ X)/\Gamma_{\text{total}}$	$\Gamma_{45}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.968 \pm 0.019 \pm \overset{+0.041}{-0.039}</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.983 \pm 0.030 \pm \overset{+0.054}{-0.051}$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(c X)/\Gamma_{\text{total}}$	$\Gamma_{46}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.234 \pm 0.012 \pm \overset{+0.018}{-0.014}</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.330 \pm 0.022 \pm \overset{+0.055}{-0.037}$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(c/\overline{c} X)/\Gamma_{\text{total}}$	$\Gamma_{47}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.202 \pm 0.023 \pm \overset{+0.053}{-0.049}</math></b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$1.313 \pm 0.037 \pm \overset{+0.088}{-0.075}$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

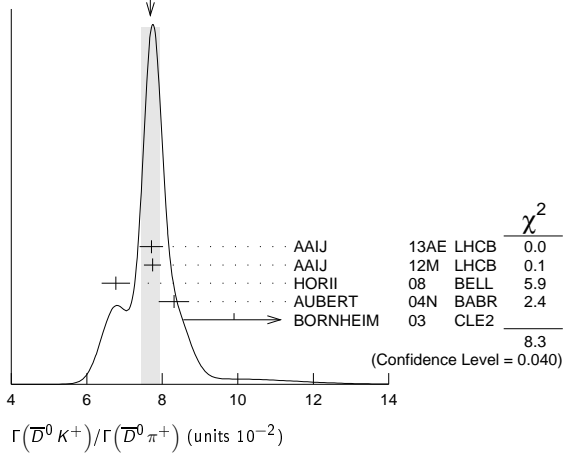
$\Gamma(\overline{D}^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{48}/\Gamma$			
VALUE (units $10^{-3}$ )	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>4.80 \pm 0.15</math> OUR FIT</b>				
<b><math>4.83 \pm 0.15</math> OUR AVERAGE</b>				
$4.90 \pm 0.07 \pm 0.22$		<sup>1</sup> AUBERT	07H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$5.0 \pm 0.6 \pm 0.3$		<sup>2</sup> ABULENCIA	06j	CDF $p\overline{p}$ at 1.96 TeV
$4.49 \pm 0.21 \pm 0.23$		<sup>3</sup> AUBERT,BE	06j	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$4.97 \pm 0.12 \pm 0.29$		<sup>1,4</sup> AHMED	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$5.0 \pm 0.7 \pm 0.6$	54	<sup>5</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$5.4 \pm \overset{+1.8}{-1.5} \pm \overset{+1.2}{-0.9}$	14	<sup>6</sup> BEBEK	87	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$4.70 \pm 0.26 \pm 0.05$		<sup>7</sup> AUBERT,B	04P	BABR Repl. by AUBERT 07H
$5.5 \pm 0.4 \pm 0.5$	304	<sup>8</sup> ALAM	94	CLE2 Repl. by AHMED 02B
$2.0 \pm 0.8 \pm 0.6$	12	<sup>5</sup> ALBRECHT	90j	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
$1.9 \pm 1.0 \pm 0.6$	7	<sup>9</sup> ALBRECHT	88k	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> ABULENCIA 06j reports $[\Gamma(B^+ \rightarrow \overline{D}^0 \pi^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \pi^+)] = 1.97 \pm 0.10 \pm 0.21$ which we multiply by our best value $B(B^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>3</sup> Uses a missing-mass method. Does not depend on $D$ branching fractions or $B^+ / B^0$ production rates.				
<sup>4</sup> AHMED 02B reports an additional uncertainty on the branching ratios to account for 4.5% uncertainty on relative production of $B^0$ and $B^+$ , which is not included here.				
<sup>5</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ and uses the MarkIII branching fractions for the $D$ .				
<sup>6</sup> BEBEK 87 value has been updated in BERKELMAN 91 to use same assumptions as noted for BORTOLETTO 92.				
<sup>7</sup> AUBERT,B 04P reports $[\Gamma(B^+ \rightarrow \overline{D}^0 \pi^+)/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^- \pi^+)] = (1.846 \pm 0.032 \pm 0.097) \times 10^{-4}$ which we divide by our best value $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>8</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ and use the CLEOII absolute $B(D^0 \rightarrow K^- \pi^+)$ and the PDG 1992 $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$ and $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .				
<sup>9</sup> ALBRECHT 88k assumes $B^0 \overline{B}^0 : B^+ B^-$ ratio is 45:55. Superseded by ALBRECHT 90j.				

$\Gamma(\overline{D}^0 \rho^+)/\Gamma_{\text{total}}$	$\Gamma_{51}/\Gamma$			
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>0.0134 \pm 0.0018</math> OUR AVERAGE</b>				
$0.0135 \pm 0.0012 \pm 0.0015$	212	<sup>1</sup> ALAM	94	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$0.013 \pm 0.004 \pm 0.004$	19	<sup>2</sup> ALBRECHT	90j	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
$0.021 \pm 0.008 \pm 0.009$	10	<sup>3</sup> ALBRECHT	88k	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ and use the CLEOII absolute $B(D^0 \rightarrow K^- \pi^+)$ and the PDG 1992 $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$ and $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ and uses the MarkIII branching fractions for the $D$ .				
<sup>3</sup> ALBRECHT 88k assumes $B^0 \overline{B}^0 : B^+ B^-$ ratio is 45:55.				



$\Gamma(\bar{D}^0 K^+)/\Gamma(\bar{D}^0 \pi^+)$  $\Gamma_{52}/\Gamma_{48}$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>7.69±0.25 OUR AVERAGE</b>	Error includes scale factor of 1.7. See the ideogram below.		
7.71±0.17±0.26	<sup>1</sup> AAIJ	13AE LHCb	$pp$ at 7 TeV
7.74±0.12±0.19	AAIJ	12M LHCb	$pp$ at 7 TeV
6.77±0.23±0.30	HORII	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
8.31±0.35±0.20	AUBERT	04N BABR	$e^+e^- \rightarrow \Upsilon(4S)$
9.9 $^{+1.4}_{-1.2} \pm 0.7_{-0.6}$	BORNHEIM	03 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
9.4 ± 0.9 ± 0.7	ABE	03D BELL	Repl. by SWAIN 03
7.7 ± 0.5 ± 0.6	SWAIN	03 BELL	Repl. by HORII 08
7.9 ± 0.9 ± 0.6	ABE	01i BELL	Repl. by ABE 03D
5.5 ± 1.4 ± 0.5	ATHANAS	98 CLE2	Repl. by BORNHEIM 03

<sup>1</sup> Uses  $B^\pm \rightarrow [K^\pm \pi^\mp \pi^\pm \pi^-]_D h^\pm$  mode.WEIGHTED AVERAGE  
7.69±0.25 (Error scaled by 1.7) $\Gamma(D_{CP(+1)} K^+)/\Gamma(D_{CP(+1)} \pi^+)$  $\Gamma_{53}/\Gamma_{49}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.087±0.007 OUR AVERAGE</b>	Error includes scale factor of 1.6.		
0.087±0.008±0.003	<sup>1,2</sup> ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.088±0.016±0.005	<sup>3</sup> AUBERT	04N BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.125±0.036±0.010	<sup>3</sup> ABE	03D BELL	Repl. by SWAIN 03
0.093±0.018±0.008	<sup>3</sup> SWAIN	03 BELL	Repl. by AUBERT 06

<sup>1</sup> Reports a double ratio of  $B(B^+ \rightarrow D_{CP(+1)} K^+)/B(B^+ \rightarrow D_{CP(+1)} \pi^+)$  and  $B(B^+ \rightarrow \bar{D}^0 K^+)/B(B^+ \rightarrow \bar{D}^0 \pi^+)$ ,  $1.13 \pm 0.16 \pm 0.08$ . We multiply by our best value of  $B(B^+ \rightarrow \bar{D}^0 K^+)/B(B^+ \rightarrow \bar{D}^0 \pi^+) = 0.083 \pm 0.006$ . Our first error is their experiment's error and the second error is systematic error from using our best value.

<sup>2</sup> ABE 06 reports  $[\Gamma(B^+ \rightarrow D_{CP(+1)} K^+)/\Gamma(B^+ \rightarrow D_{CP(+1)} \pi^+)] / [\Gamma(B^+ \rightarrow \bar{D}^0 K^+)/\Gamma(B^+ \rightarrow \bar{D}^0 \pi^+)] = 1.13 \pm 0.06 \pm 0.08$  which we multiply by our best value  $\Gamma(B^+ \rightarrow \bar{D}^0 K^+)/\Gamma(B^+ \rightarrow \bar{D}^0 \pi^+) = 0.0769 \pm 0.0025$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup>  $CP=+1$  eigenstate of  $D^0 \bar{D}^0$  system is reconstructed via  $K^+ K^-$  and  $\pi^+ \pi^-$ .

 $\Gamma(D_{CP(+1)} K^+)/\Gamma(\bar{D}^0 K^+)$  $\Gamma_{53}/\Gamma_{52}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.518±0.029 OUR AVERAGE</b>	Error includes scale factor of 1.6.		
0.504±0.019±0.006	<sup>1</sup> AAIJ	12M LHCb	$pp$ at 7 TeV
0.65 ± 0.12 ± 0.06	<sup>2</sup> AALTONEN	10A CDF	$p\bar{p}$ at 1.96 TeV
0.590±0.045±0.025	<sup>3</sup> DEL-AMO-SA...	10G BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.53 ± 0.05 ± 0.025	AUBERT	08AA BABR	Repl. by DEL-AMO-SANCHEZ 10G
0.45 ± 0.06 ± 0.02	AUBERT	06J BABR	Repl. by AUBERT 08AA

<sup>1</sup> AAIJ 12M reports  $R_{CP+} = 1.007 \pm 0.038 \pm 0.012$  which we have divided by 2.

<sup>2</sup> Reports  $R_{CP+} = 2 (B(B^- \rightarrow D_{CP(+1)} K^-) + B(B^+ \rightarrow D_{CP(+1)} K^+)) / (B(B^- \rightarrow D^0 K^-) + B(B^+ \rightarrow \bar{D}^0 K^+)) = 1.30 \pm 0.24 \pm 0.12$  that we have divided by 2.

<sup>3</sup> Reports  $R_{CP+} = 1.18 \pm 0.09 \pm 0.05$  that we have divided by 2.

 $\Gamma(D_{CP(-1)} K^+)/\Gamma(D_{CP(-1)} \pi^+)$  $\Gamma_{54}/\Gamma_{50}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.097±0.016±0.007</b>	<sup>1</sup> ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.119±0.028±0.006	<sup>2</sup> ABE	03D BELL	Repl. by SWAIN 03
0.108±0.019±0.007	<sup>2</sup> SWAIN	03 BELL	Repl. by ABE 06

<sup>1</sup> Reports a double ratio of  $B(B^+ \rightarrow D_{CP(-1)} K^+)/B(B^+ \rightarrow D_{CP(-1)} \pi^+)$  and  $B(B^+ \rightarrow \bar{D}^0 K^+)/B(B^+ \rightarrow \bar{D}^0 \pi^+)$ ,  $1.17 \pm 0.14 \pm 0.14$ . We multiply by our best value of  $B(B^+ \rightarrow \bar{D}^0 K^+)/B(B^+ \rightarrow \bar{D}^0 \pi^+) = 0.083 \pm 0.006$ . Our first error is their experiment's error and the second error is systematic error from using our best value.

<sup>2</sup>  $CP=-1$  eigenstate of  $D^0 \bar{D}^0$  system is reconstructed via  $K_S^0 \pi^0$ ,  $K_S^0 \omega$ ,  $K_S^0 \phi$ ,  $K_S^0 \eta$ , and  $K_S^0 \eta'$ .

 $\Gamma(D_{CP(-1)} K^+)/\Gamma(\bar{D}^0 K^+)$  $\Gamma_{54}/\Gamma_{52}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.54 ± 0.04 ± 0.02</b>	<sup>1</sup> DEL-AMO-SA...	10G BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.515±0.05±0.025	AUBERT	08AA BABR	Repl. by DEL-AMO-SANCHEZ 10G
0.43 ± 0.05 ± 0.02	AUBERT	06J BABR	Repl. by AUBERT 08AA

<sup>1</sup> Reports  $R_{CP+} = 1.07 \pm 0.08 \pm 0.04$  that we have divided by 2.

 $\Gamma([K^- \pi^+]_D K^+)/\Gamma_{\text{total}}$  $\Gamma_{55}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.8 × 10<sup>-7</sup></b>	90	HORII	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<6.3 × 10 <sup>-7</sup>	90	SAIGO	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma([K^- \pi^+]_D K^+)/\Gamma([K^+ \pi^-]_D K^+)$  $\Gamma_{55}/\Gamma_{56}$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>15.3±1.7 OUR AVERAGE</b>				
15.2±2.0±0.4		AAIJ	12M LHCb	$pp$ at 7 TeV
22.0±8.6±2.6		<sup>1</sup> AALTONEN	11AJ CDF	$p\bar{p}$ at 1.96 TeV
16.3 $^{+4.4}_{-4.1} \pm 0.7_{-1.3}$		HORII	11 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
11 ± 6 ± 2		DEL-AMO-SA...	10H BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7.8 $^{+6.2}_{-5.7} \pm 2.0_{-2.8}$		HORII	08 BELL	Repl. by HORII 11
<29	90	<sup>2</sup> AUBERT	05G BABR	Repl. by DEL-AMO-SANCHEZ 10H
<44	90	<sup>3</sup> SAIGO	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<26	90	<sup>4</sup> AUBERT,B	04L BABR	Repl. by AUBERT 05G

<sup>1</sup> AALTONEN 11AJ also measures the ratio separately for  $B^+ (R^+(K))$  and  $B^- (R^-(K))$  and obtains:  $R^+(K) = (42.6 \pm 13.7 \pm 2.8) \times 10^{-3}$ ,  $R^-(K) = (3.8 \pm 10.3 \pm 2.7) \times 10^{-3}$ .

<sup>2</sup> AUBERT 05G extract a constraint on the magnitude of the ratio of amplitudes  $|A(B^+ \rightarrow D^0 K^+) / A(B^+ \rightarrow \bar{D}^0 K^+)| < 0.23$  at 90% CL (Bayesian). Similar measurements from  $B^+ \rightarrow D^{*0} K^+$  are also reported.

<sup>3</sup> SAIGO 05 extract a constraint on the magnitude of the ratio of amplitudes  $|A(B^+ \rightarrow D^0 K^+) / A(B^+ \rightarrow \bar{D}^0 K^+)| < 0.27$  at 90% CL.

<sup>4</sup> AUBERT,B 04L extract a constraint on the magnitude of the ratio of amplitudes  $|A(B^+ \rightarrow D^0 K^+) / A(B^+ \rightarrow \bar{D}^0 K^+)| < 0.22$  at 90% CL.

 $\Gamma([K^- \pi^+ \pi^0]_D K^+)/\Gamma([K^+ \pi^- \pi^0]_D K^+)$  $\Gamma_{57}/\Gamma_{58}$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>16 ± 4 OUR AVERAGE</b>				
14.0±4.7±2.1		<sup>1</sup> AAIJ	15W LHCb	$pp$ at 7, 8 TeV
19.8±6.2±2.4		NAYAK	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<21	90	<sup>2</sup> LEES	11D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<39	95	<sup>3</sup> AUBERT	07BN BABR	Repl. by LEES 11D

<sup>1</sup> Uses  $D^0 \rightarrow K^- \pi^+ \pi^0$  for the favored mode, and  $D^0 \rightarrow K^+ \pi^- \pi^0$  for the suppressed mode.

<sup>2</sup> Extracts a constraint on the magnitude of the ratio of amplitudes  $|A(B^+ \rightarrow D^0 K^+) / A(B^+ \rightarrow \bar{D}^0 K^+)| < 0.13$  at 95% CL.

<sup>3</sup> Extracts a constraint on the magnitude of the ratio of amplitudes  $|A(B^+ \rightarrow D^0 K^+) / A(B^+ \rightarrow \bar{D}^0 K^+)| < 0.19$  at 95% CL.

 $\Gamma([K^- \pi^+ \pi^+ \pi^-]_D K^+)/\Gamma([K^+ \pi^- \pi^+ \pi^-]_D K^+)$  $\Gamma_{59}/\Gamma_{60}$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.24±0.27</b>	AAIJ	13AE LHCb	$pp$ at 7 TeV

 $\Gamma([K^- \pi^+]_D K^*(892)^+)/\Gamma([K^+ \pi^-]_D K^*(892)^+)$  $\Gamma_{61}/\Gamma_{62}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.066±0.031±0.010</b>	AUBERT	09AJ BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.046±0.031±0.008	AUBERT,B	05V BABR	Repl. by AUBERT 09AJ

 $\Gamma([K^- \pi^+]_D \pi^+)/\Gamma_{\text{total}}$  $\Gamma_{63}/\Gamma$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.29<math>^{+1.02}_{-0.95} \pm 0.37_{-0.48}</math></b>	HORII	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6.6 $^{+1.9}_{-1.7} \pm 0.5$	SAIGO	05 BELL	Repl. by HORII 08

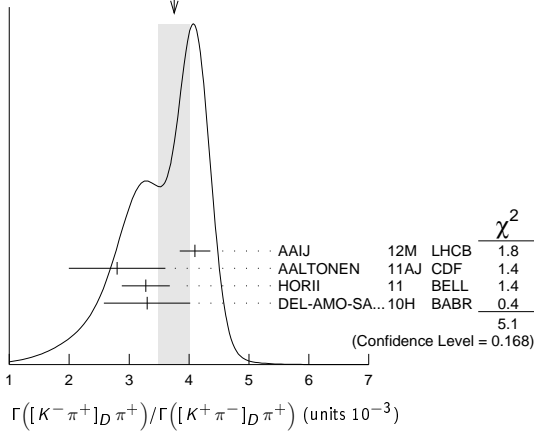
## Meson Particle Listings

 $B^\pm$  $\Gamma([K^-\pi^+]_D\pi^+)/\Gamma([K^+\pi^-]_D\pi^+)$   $\Gamma_{63}/\Gamma_{64}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.75 \pm 0.26</math> OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.		
$4.10 \pm 0.25 \pm 0.05$	AAIJ	12M LHCb	$pp$ at 7 TeV
$2.8 \pm 0.7 \pm 0.4$	<sup>1</sup> AALTONEN	11AJ CDF	$p\bar{p}$ at 1.96 TeV
$3.28^{+0.38+0.12}_{-0.36-0.18}$	HORII	11 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$3.3 \pm 0.6 \pm 0.4$	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.40^{+0.55+0.15}_{-0.53-0.22}$	HORII	08 BELL	Repl. by HORII 11
$3.5^{+1.0}_{-0.9} \pm 0.2$	SAIGO	05 BELL	Repl. by HORII 08

<sup>1</sup> AALTONEN 11AJ also measures the ratio separately for  $B^+$  ( $R^+(\pi)$ ) and  $B^-$  ( $R^-(\pi)$ ) and obtains:  $R^+(\pi) = (2.4 \pm 1.0 \pm 0.4) \times 10^{-3}$ ,  $R^-(\pi) = (3.1 \pm 1.1 \pm 0.4) \times 10^{-3}$ .

WEIGHTED AVERAGE  
 $3.75 \pm 0.26$  (Error scaled by 1.3)

 $\Gamma([K^-\pi^+\pi^0]_D\pi^+)/\Gamma([K^+\pi^-\pi^0]_D\pi^+)$   $\Gamma_{65}/\Gamma_{66}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.2 \pm 0.4</math> OUR AVERAGE</b>			
$2.35 \pm 0.49 \pm 0.06$	<sup>1</sup> AAIJ	15W LHCb	$pp$ at 7, 8 TeV
$1.89 \pm 0.54^{+0.22}_{-0.25}$	NAYAK	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $D^0 \rightarrow K^-\pi^+\pi^0$  for the favored mode, and  $D^0 \rightarrow K^+\pi^-\pi^0$  for the suppressed mode.

 $\Gamma([K^-\pi^+\pi^+\pi^-]_D\pi^+)/\Gamma([K^+\pi^-\pi^+\pi^-]_D\pi^+)$   $\Gamma_{67}/\Gamma_{68}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.7 \pm 0.4</math></b>	AAIJ	13AE LHCb	$pp$ at 7 TeV

 $\Gamma([K^-\pi^+]_{(D\pi)}\pi^+)/\Gamma([K^+\pi^-]_{(D\pi)}\pi^+)$   $\Gamma_{69}/\Gamma_{70}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.2 \pm 0.9 \pm 0.8</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma([K^-\pi^+]_{(D\gamma)}\pi^+)/\Gamma([K^+\pi^-]_{(D\gamma)}\pi^+)$   $\Gamma_{71}/\Gamma_{72}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.7 \pm 1.4 \pm 2.2</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma([K^-\pi^+]_{(D\pi)}K^+)/\Gamma([K^+\pi^-]_{(D\pi)}K^+)$   $\Gamma_{73}/\Gamma_{74}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.8 \pm 0.9 \pm 0.4</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma([K^-\pi^+]_{(D\gamma)}K^+)/\Gamma([K^+\pi^-]_{(D\gamma)}K^+)$   $\Gamma_{75}/\Gamma_{76}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.3 \pm 1.4 \pm 0.8</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma([\pi^+\pi^-\pi^0]_D K^-)/\Gamma_{\text{total}}$   $\Gamma_{77}/\Gamma$ 

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>4.6 \pm 0.8 \pm 0.4</math></b>	<sup>1</sup> AUBERT	07BJ BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$5.5 \pm 1.0 \pm 0.7$	<sup>1</sup> AUBERT,B	05T BABR	Repl. by AUBERT 07BJ

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma([K_S^0 K^+\pi^-]_D K^+)/\Gamma([K_S^0 K^+\pi^-]_D \pi^+)$   $\Gamma_{78}/\Gamma_{83}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.092 \pm 0.009 \pm 0.004</math></b>	<sup>1</sup> AAIJ	14V LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The analysis uses all of  $D \rightarrow K_S^0 K \pi$  Dalitz decays.

 $\Gamma([K_S^0 K^-\pi^+]_D K^+)/\Gamma([K_S^0 K^-\pi^+]_D \pi^+)$   $\Gamma_{79}/\Gamma_{81}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.066 \pm 0.009 \pm 0.002</math></b>	<sup>1</sup> AAIJ	14V LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The analysis uses all of  $D \rightarrow K_S^0 K \pi$  Dalitz decays.

 $\Gamma([K_S^0 K^-\pi^+]_D K^+)/\Gamma([K_S^0 K^+\pi^-]_D \pi^+)$   $\Gamma_{79}/\Gamma_{83}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.084 \pm 0.011 \pm 0.003</math></b>	<sup>1</sup> AAIJ	14V LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The Analysis uses  $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$  decays.

 $\Gamma([K^*(892)^+ K^-]_D K^+)/\Gamma([K^*(892)^- K^+]_D \pi^+)$   $\Gamma_{80}/\Gamma_{84}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.056 \pm 0.013 \pm 0.002</math></b>	<sup>1</sup> AAIJ	14V LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The Analysis uses  $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$  decays.

 $\Gamma([K^+ K^-\pi^0]_D K^+)/\Gamma([K^+ K^-\pi^0]_D \pi^+)$   $\Gamma_{85}/\Gamma_{86}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.95 \pm 0.22 \pm 0.05</math></b>	<sup>1</sup> AAIJ	15W LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Uses  $D \rightarrow K^+ K^-\pi^0$  mode.

 $\Gamma([\pi^+\pi^-\pi^0]_D K^+)/\Gamma([\pi^+\pi^-\pi^0]_D \pi^+)$   $\Gamma_{87}/\Gamma_{88}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.98 \pm 0.11 \pm 0.05</math></b>	<sup>1</sup> AAIJ	15W LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Uses  $D \rightarrow \pi^+\pi^-\pi^0$  mode.

 $\Gamma([K_S^0 K^+\pi^-]_D \pi^+)/\Gamma([K_S^0 K^+\pi^-]_D \pi^+)$   $\Gamma_{83}/\Gamma_{81}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.528 \pm 0.058 \pm 0.025</math></b>	<sup>1</sup> AAIJ	14V LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The analysis uses all of  $D \rightarrow K_S^0 K \pi$  Dalitz decays.

 $\Gamma([K^*(892)^- K^+]_D \pi^+)/\Gamma([K^*(892)^+ K^-]_D \pi^+)$   $\Gamma_{84}/\Gamma_{82}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>2.57 \pm 0.13 \pm 0.06</math></b>	<sup>1</sup> AAIJ	14V LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The Analysis uses  $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$  decays.

 $\Gamma(\bar{D}^0 K^*(892)^+)/\Gamma_{\text{total}}$   $\Gamma_{89}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.3 \pm 0.4</math> OUR AVERAGE</b>			
$5.29 \pm 0.30 \pm 0.34$	<sup>1</sup> AUBERT	06Z BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$6.1 \pm 1.6 \pm 1.7$	<sup>1</sup> MAHAPATRA	02 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$6.3 \pm 0.7 \pm 0.5$	<sup>1</sup> AUBERT	04Q BABR	Repl. by AUBERT 06Z

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_{CP(-)} K^*(892)^+)/\Gamma(\bar{D}^0 K^*(892)^+)$   $\Gamma_{90}/\Gamma_{89}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.515 \pm 0.135 \pm 0.065</math></b>	<sup>1</sup> AUBERT	09AJ BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.325 \pm 0.13 \pm 0.04$	<sup>2</sup> AUBERT,B	05U BABR	Repl. by AUBERT 09AJ

<sup>1</sup> The authors report  $R_{CP-} = 1.03 \pm 0.27 \pm 0.13$  which is, assuming  $CP$  conservation, twice the value of the quoted above branching ratio,  
<sup>2</sup> The authors report  $R_{CP-} = 0.65 \pm 0.26 \pm 0.08$  which is, assuming  $CP$  conservation, twice the value of the quoted above branching ratio.

 $\Gamma(D_{CP(+)} K^*(892)^+)/\Gamma(\bar{D}^0 K^*(892)^+)$   $\Gamma_{91}/\Gamma_{89}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.085 \pm 0.175 \pm 0.045</math></b>	<sup>1</sup> AUBERT	09AJ BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.98 \pm 0.20 \pm 0.055$	<sup>2</sup> AUBERT,B	05U BABR	Repl. by AUBERT 09AJ

<sup>1</sup> The authors report  $R_{CP+} = 2.17 \pm 0.35 \pm 0.09$  which is, assuming  $CP$  conservation, twice the value of the quoted above branching ratio,  
<sup>2</sup> The authors report  $R_{CP+} = 1.96 \pm 0.40 \pm 0.11$  which is, assuming  $CP$  conservation, twice the value of the quoted above branching ratio.

 $\Gamma(\bar{D}^0 K^+\pi^-\pi^-)/\Gamma(\bar{D}^0 \pi^+\pi^+\pi^-)$   $\Gamma_{92}/\Gamma_{98}$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>9.4 \pm 1.3 \pm 0.9</math></b>	AAIJ	12T LHCb	$pp$ at 7 TeV

 $\Gamma(D_{CP(+)} K^+\pi^-\pi^+)/\Gamma([K^+\pi^-]_D K^+\pi^-\pi^+)$   $\Gamma_{95}/\Gamma_{93}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.040 \pm 0.064</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $\Gamma([K^-\pi^+]_D K^+\pi^-\pi^+)/\Gamma([K^+\pi^-]_D K^+\pi^-\pi^+)$   $\Gamma_{94}/\Gamma_{93}$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>85^{+36}_{-33}</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

See key on page 601

## Meson Particle Listings

 $B^\pm$ 

$\Gamma(\bar{D}^0 K^+ \bar{K}^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{96}/\Gamma$
VALUE (units $10^{-4}$ )				
<b><math>5.5 \pm 1.4 \pm 0.8</math></b>	<sup>1</sup> DRUTSKOY 02	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0 K^+ \bar{K}^*(892)^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{97}/\Gamma$
VALUE (units $10^{-4}$ )				
<b><math>7.5 \pm 1.3 \pm 1.1</math></b>	<sup>1</sup> DRUTSKOY 02	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{98}/\Gamma$
VALUE				
<b><math>0.0057 \pm 0.0022</math> OUR FIT</b>	Error includes scale factor of 3.6.			
<b><math>0.0115 \pm 0.0029 \pm 0.0021</math></b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)/\Gamma(\bar{D}^0 \pi^+)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{98}/\Gamma_{48}$
VALUE				
<b><math>1.2 \pm 0.4</math> OUR FIT</b>	Error includes scale factor of 3.8.			
<b><math>1.27 \pm 0.06 \pm 0.11</math></b>	AAIJ	11E	LHCB $pp$ at 7 TeV	

$\Gamma([K^- \pi^+]_D \pi^+ \pi^- \pi^+)/\Gamma([K^+ \pi^-]_D K^+ \pi^- \pi^+)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{99}/\Gamma_{93}$
VALUE (units $10^{-4}$ )				
<b><math>42.7 \pm 5.6</math></b>	AAIJ	15Bc	LHCB $pp$ at 7, 8 TeV	

$\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^- \text{ nonresonant})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{100}/\Gamma$
VALUE				
<b><math>0.0051 \pm 0.0034 \pm 0.0023</math></b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(\bar{D}^0 \pi^+ \rho^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{101}/\Gamma$
VALUE				
<b><math>0.0045 \pm 0.0023 \pm 0.0020</math></b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(\bar{D}^0 a_1(1260)^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{102}/\Gamma$
VALUE				
<b><math>0.0045 \pm 0.0019 \pm 0.0031</math></b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(\bar{D}^0 \omega \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{103}/\Gamma$
VALUE				
<b><math>0.0041 \pm 0.0007 \pm 0.0006</math></b>	<sup>1</sup> ALEXANDER 01B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The signal is consistent with all observed  $\omega \pi^+$  having proceeded through the  $\rho^+$  resonance at mass  $1349 \pm 25^{+10}_{-5}$  MeV and width  $547 \pm 86^{+46}_{-45}$  MeV.

$\Gamma(D^*(2010)^- \pi^+ \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{104}/\Gamma$
VALUE (units $10^{-3}$ )	CL% EVTS			
<b><math>1.35 \pm 0.22</math> OUR AVERAGE</b>				

1.25  $\pm 0.08 \pm 0.22$  <sup>1</sup> ABE 04D BELL  $e^+ e^- \rightarrow \Upsilon(4S)$   
 1.9  $\pm 0.7 \pm 0.3$  14 <sup>2</sup> ALAM 94 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$   
 2.6  $\pm 1.4 \pm 0.7$  11 <sup>3</sup> ALBRECHT 90J ARG  $e^+ e^- \rightarrow \Upsilon(4S)$   
 2.4  $\pm 1.7 \pm 1.0$  3 <sup>4</sup> BEBEK 87 CLEO  $e^+ e^- \rightarrow \Upsilon(4S)$   
 -1.6 -0.6

• • • We do not use the following data for averages, fits, limits, etc. • • •

<4. 90 <sup>5</sup> BORTOLETTO92 CLEO  $e^+ e^- \rightarrow \Upsilon(4S)$   
 5.  $\pm 2.$   $\pm 3.$  7 <sup>6</sup> ALBRECHT 87c ARG  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II  $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses the Mark III branching fractions for the  $D$ .<sup>4</sup> BEBEK 87 value has been updated in BERKELMAN 91 to use same assumptions as noted for BORTOLETTO 92.<sup>5</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$  and  $D^*(2010)$ . The authors also find the product branching fraction into  $D^{**} \pi$  followed by  $D^{**} \rightarrow D^*(2010) \pi$  to be  $0.0014^{+0.0008}_{-0.0006} \pm 0.0003$  where  $D^{**}$  represents all orbitally excited  $D$  mesons.<sup>6</sup> ALBRECHT 87c use PDG 86 branching ratios for  $D$  and  $D^*(2010)$  and assume  $B(\Upsilon(4S) \rightarrow B^+ B^-) = 55\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = 45\%$ . Superseded by ALBRECHT 90J.

$\Gamma(\bar{D}_1(2420)^0 \pi^+, \bar{D}_1^0 \rightarrow D^*(2010)^- \pi^+)/\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{105}/\Gamma_{98}$
VALUE (units $10^{-2}$ )				
<b><math>9.3 \pm 1.6 \pm 0.9</math></b>	<sup>1</sup> AAIJ	11E	LHCB $pp$ at 7 TeV	

<sup>1</sup> AAIJ 11E reports  $(9.3 \pm 1.6 \pm 0.9) \times 10^{-2}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}_1(2420)^0 \pi^+, \bar{D}_1^0 \rightarrow D^*(2010)^- \pi^+)/\Gamma(B^+ \rightarrow \bar{D}^0 \pi^+ \pi^+ \pi^-)] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ .

$\Gamma(D^- \pi^+ \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{106}/\Gamma$
VALUE (units $10^{-3}$ )	CL% EVTS			
<b><math>1.07 \pm 0.05</math> OUR AVERAGE</b>				

1.08  $\pm 0.03 \pm 0.05$  <sup>1</sup> AUBERT 09AB BABR  $e^+ e^- \rightarrow \Upsilon(4S)$   
 1.02  $\pm 0.04 \pm 0.15$  <sup>1</sup> ABE 04D BELL  $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.4 90 <sup>2</sup> ALAM 94 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$   
 <7 90 <sup>3</sup> BORTOLETTO92 CLEO  $e^+ e^- \rightarrow \Upsilon(4S)$

2.5  $\pm 4.1 \pm 2.4$  1 <sup>4</sup> BEBEK 87 CLEO  $e^+ e^- \rightarrow \Upsilon(4S)$   
 -2.3 -0.8

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the Mark III  $B(D^+ \rightarrow K^- 2\pi^+)$ .<sup>3</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ . The product branching fraction into  $D_6^*(2340) \pi$  followed by  $D_6^*(2340) \rightarrow D \pi$  is  $< 0.005$  at 90% CL and into  $D_2^*(2460)$  followed by  $D_2^*(2460) \rightarrow D \pi$  is  $< 0.004$  at 90% CL.<sup>4</sup> BEBEK 87 assume the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ .  $B(D^- \rightarrow K^+ \pi^- \pi^-) = (9.1 \pm 1.3 \pm 0.4)\%$  is assumed.

$\Gamma(D^- K^+ \pi^+)/\Gamma(D^- \pi^+ \pi^+)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{107}/\Gamma_{106}$
VALUE (units $10^{-2}$ )				
<b><math>7.20 \pm 0.19 \pm 0.21</math></b>	AAIJ	15V	LHCB $pp$ at 7, 8 TeV	

$\Gamma(D_0^*(2400)^0 K^+, D_0^{*0} \rightarrow D^- \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{108}/\Gamma$
VALUE (units $10^{-4}$ )				
<b><math>6.1 \pm 1.9 \pm 1.5</math></b>	<sup>1</sup> AAIJ	15V	LHCB $pp$ at 7, 8 TeV	

<sup>1</sup> Performs the amplitude analysis by fitting the square-Dalitz-plot distribution.

$\Gamma(D_1^*(2760)^0 K^+, D_1^{*0} \rightarrow D^- \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{109}/\Gamma$
VALUE (units $10^{-4}$ )				
<b><math>3.6 \pm 0.9 \pm 0.8</math></b>	<sup>1</sup> AAIJ	15V	LHCB $pp$ at 7, 8 TeV	

<sup>1</sup> Performs the amplitude analysis by fitting the square-Dalitz-plot distribution.

$\Gamma(D_2^*(2460)^0 K^+, D_2^{*0} \rightarrow D^- \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{110}/\Gamma$
VALUE (units $10^{-4}$ )				
<b><math>23.2 \pm 1.1 \pm 2.0</math></b>	<sup>1</sup> AAIJ	15V	LHCB $pp$ at 7, 8 TeV	

<sup>1</sup> Performs the amplitude analysis by fitting the square-Dalitz-plot distribution.

$\Gamma(D^+ K^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{111}/\Gamma$
VALUE (units $10^{-6}$ )	CL%			
<b>&lt;2.9</b>	90	<sup>1</sup> DEL-AMO-SA...10K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<5.0 90 <sup>1</sup> AUBERT,B 05E BABR Repl. by DEL-AMO-SANCHEZ 10K

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^+ K^{*0})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{112}/\Gamma$
VALUE (units $10^{-6}$ )	CL%			
<b>&lt;1.8</b>	90	AAIJ	13R LHCB $pp$ at 7 TeV	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<3.0 90 <sup>1</sup> DEL-AMO-SA...10K BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^+ \bar{K}^{*0})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{113}/\Gamma$
VALUE (units $10^{-6}$ )	CL%			
<b>&lt;1.4</b>	90	AAIJ	13R LHCB $pp$ at 7 TeV	

$\Gamma(\bar{D}^*(2007)^0 \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{114}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS			
<b><math>5.18 \pm 0.26</math> OUR AVERAGE</b>				

5.52  $\pm 0.17 \pm 0.42$  <sup>1</sup> AUBERT 07H BABR  $e^+ e^- \rightarrow \Upsilon(4S)$   
 5.5  $\pm 0.4 \pm 0.2$  2,3 AUBERT,BE 06J BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

4.34  $\pm 0.47 \pm 0.18$  <sup>4</sup> BRANDENB... 98 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$   
 5.2  $\pm 0.7 \pm 0.7$  71 <sup>5</sup> ALAM 94 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$

7.2  $\pm 1.8 \pm 1.6$  <sup>6</sup> BORTOLETTO92 CLEO  $e^+ e^- \rightarrow \Upsilon(4S)$   
 4.0  $\pm 1.4 \pm 1.2$  9 <sup>6</sup> ALBRECHT 90J ARG  $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.7  $\pm 4.4$  <sup>7</sup> BEBEK 87 CLEO  $e^+ e^- \rightarrow \Upsilon(4S)$

# Meson Particle Listings

$B^\pm$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AUBERT,BE 06j reports  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 \pi^+)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^0 \pi^+)] = 1.14 \pm 0.07 \pm 0.04$  which we multiply by our best value  $B(B^+ \rightarrow \bar{D}^0 \pi^+) = (4.80 \pm 0.15) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> Uses a missing-mass method. Does not depend on  $D$  branching fractions or  $B^+/B^0$  production rates.  
<sup>4</sup> BRANDENBURG 98 assume equal production of  $B^+$  and  $B^0$  at  $\Upsilon(4S)$  and use the  $D^*$  reconstruction technique. The first error is their experiment's error and the second error is the systematic error from the PDG 96 value of  $B(D^* \rightarrow D\pi)$ .  
<sup>5</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEOII  $B(D^*(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .  
<sup>6</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses MarkIII branching fractions for the  $D$  and  $D^*(2010)$ .  
<sup>7</sup> This is a derived branching ratio, using the inclusive pion spectrum and other two-body  $B$  decays. BEBEK 87 assume the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ .

$\Gamma(\bar{D}^*(2007)^0 \omega \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{117}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0045 ± 0.0010 ± 0.0007</b>	<sup>1</sup> ALEXANDER 01B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The signal is consistent with all observed  $\omega \pi^+$  having proceeded through the  $\rho^+$  resonance at mass  $1349 \pm 25^{+10}_{-5}$  MeV and width  $547 \pm 86^{+46}_{-45}$  MeV.

$\Gamma(\bar{D}^*(2007)^0 \rho^+)/\Gamma_{\text{total}}$   $\Gamma_{118}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0098 ± 0.0017 OUR AVERAGE</b>				

0.0098 ± 0.0006 ± 0.0017		<sup>1</sup> CSORNA 03	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
0.010 ± 0.006 ± 0.004	7	<sup>2</sup> ALBRECHT 90j	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0168 ± 0.0021 ± 0.0028	86	<sup>3</sup> ALAM 94	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$  resonance. The second error combines the systematic and theoretical uncertainties in quadrature. CSORNA 03 includes data used in ALAM 94. A full angular fit to three complex helicity amplitudes is performed.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses MarkIII branching fractions for the  $D$  and  $D^*(2010)$ .  
<sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEOII  $B(D^*(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ . The nonresonant  $\pi^+ \pi^0$  contribution under the  $\rho^+$  is negligible.

$\Gamma(\bar{D}^*(2007)^0 K^+)/\Gamma_{\text{total}}$   $\Gamma_{119}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.20 ± 0.34 OUR AVERAGE</b>			

4.21 $\pm 0.30$ $\pm 0.26$ ± 0.21	<sup>1</sup> AUBERT 05N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
4.0 ± 1.1 ± 0.2	<sup>2</sup> ABE 01i	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> AUBERT 05N reports  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 K^+)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^*(2007)^0 \pi^+)] = 0.0813 \pm 0.0040^{+0.0042}_{-0.0031}$  which we multiply by our best value  $B(B^+ \rightarrow \bar{D}^*(2007)^0 \pi^+) = (5.18 \pm 0.26) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> ABE 01i reports  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 K^+)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^*(2007)^0 \pi^+)] = 0.078 \pm 0.019 \pm 0.009$  which we multiply by our best value  $B(B^+ \rightarrow \bar{D}^*(2007)^0 \pi^+) = (5.18 \pm 0.26) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{D}_{CP(1)}^{*0} K^+)/\Gamma_{\text{total}}$   $\Gamma_{120}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.75 ± 0.29 ± 0.23 ± 0.22</b>	<sup>1</sup> AUBERT 08BF	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> AUBERT 08BF reports  $[\Gamma(B^+ \rightarrow \bar{D}_{CP(1)}^{*0} K^+)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^*(2007)^0 K^+)] = 0.655 \pm 0.065 \pm 0.020$  which we multiply by our best value  $B(B^+ \rightarrow \bar{D}^*(2007)^0 K^+) = (4.20 \pm 0.34) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{D}_{CP(1)}^{*0} K^+)/\Gamma(\bar{D}_{CP(1)}^{*0} \pi^+)$   $\Gamma_{120}/\Gamma_{115}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.095 ± 0.017 OUR AVERAGE</b>			

0.11 ± 0.02 ± 0.02	<sup>1</sup> ABE 06	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
0.086 ± 0.021 ± 0.007	<sup>2</sup> AUBERT 05N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Reports a double ratio of  $B(B^+ \rightarrow D_{CP(1)}^{*0} K^+)/B(B^+ \rightarrow D_{CP(1)}^{*0} \pi^+)$  and  $B(B^+ \rightarrow \bar{D}^{*0} K^+)/B(B^+ \rightarrow \bar{D}^{*0} \pi^+)$ ,  $1.41 \pm 0.25 \pm 0.06$ . We multiply by our best value of  $B(B^+ \rightarrow \bar{D}^{*0} K^+)/B(B^+ \rightarrow \bar{D}^{*0} \pi^+) = 0.080 \pm 0.011$ . Our first error is their experiment's error and the second error is systematic error from using our best value.  
<sup>2</sup> Uses  $D^{*0} \rightarrow D^0 \pi^0$  with  $D^0$  reconstructed in the  $CP$ -even eigenstates  $K^+ K^-$  and  $\pi^+ \pi^-$ .

$\Gamma(\bar{D}_{CP(-1)}^{*0} K^+)/\Gamma_{\text{total}}$   $\Gamma_{121}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.31 ± 0.27 ± 0.20 ± 0.18</b>	<sup>1</sup> AUBERT 08BF	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> AUBERT 08BF reports  $[\Gamma(B^+ \rightarrow \bar{D}_{CP(-1)}^{*0} K^+)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^*(2007)^0 K^+)] = 0.55 \pm 0.06 \pm 0.02$  which we multiply by our best value  $B(B^+ \rightarrow \bar{D}^*(2007)^0 K^+) = (4.20 \pm 0.34) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{D}_{CP(-1)}^{*0} K^+)/\Gamma(D_{CP(-1)}^{*0} \pi^+)$   $\Gamma_{121}/\Gamma_{116}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.09 ± 0.03 ± 0.01</b>	<sup>1</sup> ABE 06	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Reports a double ratio of  $B(B^+ \rightarrow (D_{CP(-1)}^{*0})^0 K^+)/B(B^+ \rightarrow (D_{CP(-1)}^{*0})^0 \pi^+)$  and  $B(B^+ \rightarrow \bar{D}^{*0} K^+)/B(B^+ \rightarrow \bar{D}^{*0} \pi^+)$ ,  $1.15 \pm 0.31 \pm 0.12$ . We multiply by our best value of  $B(B^+ \rightarrow \bar{D}^{*0} K^+)/B(B^+ \rightarrow \bar{D}^{*0} \pi^+) = 0.080 \pm 0.011$ . Our first error is their experiment's error and the second error is systematic error from using our best value.

$\Gamma(\bar{D}^*(2007)^0 K^*(892)^+)/\Gamma_{\text{total}}$   $\Gamma_{122}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>8.1 ± 1.4 OUR AVERAGE</b>			

8.3 ± 1.1 ± 1.0	<sup>1</sup> AUBERT 04k	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
7.2 ± 2.2 ± 2.6	<sup>2</sup> MAHAPATRA 02	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and an unpolarized final state.

$\Gamma(\bar{D}^*(2007)^0 K^+ \bar{K}^0)/\Gamma_{\text{total}}$   $\Gamma_{123}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;10.6</b>	90	<sup>1</sup> DRUTSKOY 02	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^*(2007)^0 K^+ K^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{124}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>15.3 ± 3.1 ± 2.9</b>	<sup>1</sup> DRUTSKOY 02	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^*(2007)^0 \pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{125}/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.03 ± 0.12 OUR AVERAGE</b>				

1.055 ± 0.047 ± 0.129		<sup>1</sup> MAJUMDER 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
0.94 ± 0.20 ± 0.17	48	<sup>2,3</sup> ALAM 94	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEOII  $B(D^*(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .  
<sup>3</sup> The three pion mass is required to be between 1.0 and 1.6 GeV consistent with an  $a_1$  meson. (If this channel is dominated by  $a_1^+$ , the branching ratio for  $\bar{D}^{*0} a_1^+$  is twice that for  $\bar{D}^{*0} \pi^+ \pi^+ \pi^-$ .)

$\Gamma(\bar{D}^*(2007)^0 a_1(1260)^+)/\Gamma_{\text{total}}$   $\Gamma_{126}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0188 ± 0.0040 ± 0.0034</b>	<sup>1,2</sup> ALAM 94	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> ALAM 94 value is twice their  $\Gamma(\bar{D}^*(2007)^0 \pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$  value based on their observation that the three pions are dominantly in the  $a_1(1260)$  mass range 1.0 to 1.6 GeV.  
<sup>2</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEOII  $B(D^*(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

$\Gamma(\bar{D}^*(2007)^0 \pi^- \pi^+ \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{127}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0180 ± 0.0024 ± 0.0027</b>	<sup>1</sup> ALEXANDER 01B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The signal is consistent with all observed  $\omega \pi^+$  having proceeded through the  $\rho^+$  resonance at mass  $1349 \pm 25^{+10}_{-5}$  MeV and width  $547 \pm 86^{+46}_{-45}$  MeV.

$\Gamma(\bar{D}^{*0} 3\pi^+ 2\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{128}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>5.67 ± 0.91 ± 0.85</b>	<sup>1</sup> MAJUMDER 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^*(2010)^+ \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{129}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.6 × 10<sup>-6</sup></b>		<sup>1</sup> IWABUCHI 08	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

- • • We do not use the following data for averages, fits, limits, etc. • • •  
 <1.7 × 10<sup>-4</sup> 90 <sup>2</sup> BRANDENB... 98 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> BRANDENBURG 98 assume equal production of  $B^+$  and  $B^0$  at  $\Upsilon(4S)$  and use the  $D^*$  partial reconstruction technique. The first error is their experiment's error and the second error is the systematic error from the PDG 96 value of  $B(D^* \rightarrow D\pi)$ .

$\Gamma(D^*(2010)^+ K^0)/\Gamma_{\text{total}}$					$\Gamma_{130}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<9.0 \times 10^{-6}$	90	<sup>1</sup> AUBERT,B	05E BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<9.5 \times 10^{-5}$	90	<sup>1</sup> GRITSAN	01 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{131}/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT	
$0.0152 \pm 0.0071 \pm 0.0001$	26	<sup>1</sup> ALBRECHT	90J ARG	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.043 \pm 0.013 \pm 0.026$	24	<sup>2</sup> ALBRECHT	87C ARG	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> ALBRECHT 90J reports $0.018 \pm 0.007 \pm 0.005$ from a measurement of $[\Gamma(B^+ \rightarrow D^*(2010)^- \pi^+ \pi^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$ assuming $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and uses MarkIII branching fractions for the $D$ .					
<sup>2</sup> ALBRECHT 87C use PDG 86 branching ratios for $D$ and $D^*(2010)$ and assume $B(\gamma(4S) \rightarrow B^+ B^-) = 55\%$ and $B(\gamma(4S) \rightarrow B^0 \bar{B}^0) = 45\%$ . Superseded by ALBRECHT 90J.					

$\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{132}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
$2.56 \pm 0.26 \pm 0.33$		<sup>1</sup> MAJUMDER	04 BELL	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<10$	90	<sup>2</sup> ALBRECHT	90J ARG	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and uses MarkIII branching fractions for the $D$ and $D^*(2010)$ .					

$\Gamma(D^{**0} \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{133}/\Gamma$
$D^{**0}$ represents an excited state with mass $2.2 < M < 2.8$ GeV/ $c^2$ .					
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
$5.9 \pm 1.3 \pm 0.2$		<sup>1,2</sup> AUBERT,BE	06J BABR	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> AUBERT,BE 06J reports $[\Gamma(B^+ \rightarrow D^{**0} \pi^+)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow D^0 \pi^+)] = 1.22 \pm 0.13 \pm 0.23$ which we multiply by our best value $B(B^+ \rightarrow D^0 \pi^+) = (4.80 \pm 0.15) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Uses a missing-mass method. Does not depend on $D$ branching fractions or $B^+ B^0$ production rates.					

$\Gamma(D_s^*(2420)^0 \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{134}/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT	
$0.0015 \pm 0.0006$ OUR AVERAGE		Error includes scale factor of 1.3.			
$0.0011 \pm 0.0005 \pm 0.0002$	8	<sup>1</sup> ALAM	94 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
$0.0025 \pm 0.0007 \pm 0.0006$		<sup>2</sup> ALBRECHT	94D ARG	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the CLEOII $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$ and absolute $B(D^0 \rightarrow K^- \pi^+)$ and the PDG 1992 $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$ and assuming $B(D_s(2420)^0 \rightarrow D^*(2010)^+ \pi^-) = 67\%$ .					
<sup>2</sup> ALBRECHT 94D assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the CLEOII $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$ assuming $B(D_s(2420)^0 \rightarrow D^*(2010)^+ \pi^-) = 67\%$ .					

$\Gamma(D_s(2420)^0 \pi^+ \times B(\bar{D}_1^0 \rightarrow \bar{D}^0 \pi^+ \pi^-))/\Gamma_{\text{total}}$					$\Gamma_{135}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
$2.5 \substack{+1.6 \\ -1.4}$ OUR FIT		Error includes scale factor of 4.0.			
$1.85 \pm 0.29 \substack{+0.35 \\ -0.55}$		<sup>1</sup> ABE	05A BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_1(2420)^0 \pi^+ \times B(\bar{D}_1^0 \rightarrow \bar{D}^0 \pi^+ \pi^-))/\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)$					$\Gamma_{135}/\Gamma_{98}$
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
$4.4 \substack{+3.3 \\ -2.6}$ OUR FIT		Error includes scale factor of 4.0.			
$10.3 \pm 1.5 \pm 0.9$		AAIJ	11E LHCB	$pp$ at 7 TeV	
$\Gamma(\bar{D}_1(2420)^0 \pi^+ \times B(\bar{D}_1^0 \rightarrow \bar{D}^0 \pi^+ \pi^- (\text{nonresonant}))/\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)$					$\Gamma_{136}/\Gamma_{98}$
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
$4.0 \pm 0.7 \pm 0.5$		<sup>1</sup> AAII	11E LHCB	$pp$ at 7 TeV	
<sup>1</sup> Excludes decays where $\bar{D}_1(2420)^0 \rightarrow D^*(2010)^- \pi^+$ .					

$\Gamma(\bar{D}_s^*(2462)^0 \pi^+ \times B(\bar{D}_s^*(2462)^0 \rightarrow D^- \pi^+))/\Gamma_{\text{total}}$					$\Gamma_{137}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
$3.5 \pm 0.4$ OUR AVERAGE					
$3.5 \pm 0.2 \pm 0.4$		<sup>1</sup> AUBERT	09AB BABR	$e^+e^- \rightarrow \gamma(4S)$	
$3.4 \pm 0.3 \pm 0.72$		<sup>1</sup> ABE	04D BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_s^*(2462)^0 \pi^+ \times B(\bar{D}_s^0 \rightarrow \bar{D}^0 \pi^- \pi^+))/\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)$					$\Gamma_{138}/\Gamma_{98}$
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
$4.0 \pm 1.0 \pm 0.4$		AAIJ	11E LHCB	$pp$ at 7 TeV	

$\Gamma(\bar{D}_s^*(2462)^0 \pi^+ \times B(\bar{D}_s^0 \rightarrow \bar{D}^0 \pi^- \pi^+ (\text{nonresonant}))/\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)$					$\Gamma_{139}/\Gamma_{98}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.0 \times 10^{-2}$	90	<sup>1</sup> AAII	11E LHCB	$pp$ at 7 TeV	
<sup>1</sup> Excludes decays where $\bar{D}_s^*(2462)^0 \rightarrow D^*(2010)^- \pi^+$ .					

$\Gamma(\bar{D}_s^*(2462)^0 \pi^+ \times B(\bar{D}_s^0 \rightarrow D^*(2010)^- \pi^+))/\Gamma(\bar{D}^0 \pi^+ \pi^+ \pi^-)$					$\Gamma_{140}/\Gamma_{98}$
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
$3.9 \pm 1.2 \pm 0.4$		<sup>1</sup> AAII	11E LHCB	$pp$ at 7 TeV	
<sup>1</sup> Uses $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5)\%$ .					

$\Gamma(\bar{D}_s^*(2400)^0 \pi^+ \times B(\bar{D}_s^0(2400)^0 \rightarrow D^- \pi^+))/\Gamma_{\text{total}}$					$\Gamma_{141}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
$6.4 \pm 1.4$ OUR AVERAGE					
$6.8 \pm 0.3 \pm 2.0$		<sup>1</sup> AUBERT	09AB BABR	$e^+e^- \rightarrow \gamma(4S)$	
$6.1 \pm 0.6 \pm 1.8$		<sup>1</sup> ABE	04D BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_1(2421)^0 \pi^+ \times B(\bar{D}_1(2421)^0 \rightarrow D^{*-} \pi^+))/\Gamma_{\text{total}}$					$\Gamma_{142}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
$6.8 \pm 0.7 \pm 1.3$		<sup>1</sup> ABE	04D BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_s^*(2462)^0 \pi^+ \times B(\bar{D}_s^*(2462)^0 \rightarrow D^{*-} \pi^+))/\Gamma_{\text{total}}$					$\Gamma_{143}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
$1.8 \pm 0.3 \pm 0.4$		<sup>1</sup> ABE	04D BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_1(2427)^0 \pi^+ \times B(\bar{D}_1^0(2427)^0 \rightarrow D^{*-} \pi^+))/\Gamma_{\text{total}}$					$\Gamma_{144}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
$5.0 \pm 0.4 \pm 1.1$		<sup>1</sup> ABE	04D BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_1(2420)^0 \pi^+ \times B(\bar{D}_1^0 \rightarrow \bar{D}^{*0} \pi^+ \pi^-))/\Gamma_{\text{total}}$					$\Gamma_{145}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.06$	90	<sup>1</sup> ABE	05A BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_1^*(2420)^0 \rho^+)/\Gamma_{\text{total}}$					$\Gamma_{146}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.0014$	90	<sup>1</sup> ALAM	94 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the CLEOII $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$ assuming $B(D_s(2420)^0 \rightarrow D^*(2010)^+ \pi^-) = 67\%$ .					

$\Gamma(\bar{D}_s^*(2460)^0 \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{147}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.0013$	90	<sup>1</sup> ALAM	94 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<0.0028$	90	<sup>2</sup> ALAM	94 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
$<0.0023$	90	<sup>3</sup> ALBRECHT	94D ARG	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the MarkIII $B(D^+ \rightarrow K^- 2\pi^+)$ and $B(D_s^*(2460)^0 \rightarrow D^+ \pi^-) = 30\%$ .					
<sup>2</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the MarkIII $B(D^+ \rightarrow K^- 2\pi^+)$ , the CLEOII $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$ and $B(D_s^*(2460)^0 \rightarrow D^*(2010)^+ \pi^-) = 20\%$ .					
<sup>3</sup> ALBRECHT 94D assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the CLEOII $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$ and $B(D_s^*(2460)^0 \rightarrow D^*(2010)^+ \pi^-) = 30\%$ .					

$\Gamma(\bar{D}_s^*(2460)^0 \pi^+ \times B(\bar{D}_s^0 \rightarrow \bar{D}^{*0} \pi^+ \pi^-))/\Gamma_{\text{total}}$					$\Gamma_{148}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.22$	90	<sup>1</sup> ABE	05A BELL	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\bar{D}_s^*(2460)^0 \rho^+)/\Gamma_{\text{total}}$					$\Gamma_{149}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.0047$	90	<sup>1</sup> ALAM	94 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
$<0.005$	90	<sup>2</sup> ALAM	94 CLE2	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the MarkIII $B(D^+ \rightarrow K^- 2\pi^+)$ and $B(D_s^*(2460)^0 \rightarrow D^+ \pi^-) = 30\%$ .					
<sup>2</sup> ALAM 94 assume equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ and use the MarkIII $B(D^+ \rightarrow K^- 2\pi^+)$ , the CLEOII $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$ and $B(D_s^*(2460)^0 \rightarrow D^*(2010)^+ \pi^-) = 20\%$ .					

# Meson Particle Listings

$B^\pm$

$\Gamma(D^0 D_s^+)/\Gamma_{\text{total}}$				$\Gamma_{150}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>9.0 \pm 0.9</math> OUR AVERAGE</b>				
$8.6 \pm 0.2 \pm 1.1$	<sup>1</sup> AAIJ	13AP	LHCB $p\bar{p}$ at 7 TeV	
$9.5 \pm 2.0 \pm 0.8$	<sup>2</sup> AUBERT	06N	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$9.8 \pm 2.6 \pm 0.9$	<sup>3</sup> GIBAUT	96	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
$14 \pm 8 \pm 1$	<sup>4</sup> ALBRECHT	92G	ARG $e^+e^- \rightarrow \Upsilon(4S)$	
$13 \pm 6 \pm 1$	<sup>5</sup> BORTOLETTO	090	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses $B(B^0 \rightarrow D^- D_s^+) = (7.2 \pm 0.8) \times 10^{-3}$ .				
<sup>2</sup> AUBERT 06N reports $(0.92 \pm 0.14 \pm 0.18) \times 10^{-2}$ from a measurement of $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.0462 \pm 0.0062$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>3</sup> GIBAUT 96 reports $0.0126 \pm 0.0022 \pm 0.0025$ from a measurement of $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.035$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>4</sup> ALBRECHT 92G reports $0.024 \pm 0.012 \pm 0.004$ from a measurement of $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990 $D^0$ branching ratios, e.g., $B(D^0 \rightarrow K^-\pi^+) = 3.71 \pm 0.25\%$ .				
<sup>5</sup> BORTOLETTO 90 reports $0.029 \pm 0.013$ from a measurement of $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.02$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(D_{s0}^*(2317)^+ \bar{D}^0, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{151}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.79 \pm 0.15</math> OUR AVERAGE</b>				
$0.79 \pm 0.17 \pm 0.02$	<sup>1,2</sup> CHOI	15A	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
$0.80 \pm 0.35 \pm 0.07$	<sup>2,3</sup> AUBERT,B	04s	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$0.65 \pm 0.26 \pm 0.06$	<sup>2,4</sup> KROKOVNY	03B	BELL Repl. by CHOI 15A	
<sup>1</sup> CHOI 15A reports $(8.0 \pm 1.3 \pm 1.1 \pm 0.4) \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_{s0}^*(2317)^+ \bar{D}^0, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow K^+K^-\pi^+)]$ assuming $B(D_s^+ \rightarrow K^+K^-\pi^+) = (5.39 \pm 0.21) \times 10^{-2}$ , which we rescale to our best value $B(D_s^+ \rightarrow K^+K^-\pi^+) = (5.45 \pm 0.17) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>3</sup> AUBERT,B 04s reports $(1.0 \pm 0.3 \pm 0.4) \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_{s0}^*(2317)^+ \bar{D}^0, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>4</sup> KROKOVNY 03B reports $(0.81 \pm 0.30 \pm 0.24) \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_{s0}^*(2317)^+ \bar{D}^0, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(D_{s0}(2317)^+ \bar{D}^0 \times B(D_{s0}(2317)^+ \rightarrow D_s^{*+} \gamma))/\Gamma_{\text{total}}$				$\Gamma_{152}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.76</b>	90	<sup>1</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_{s0}(2317)^+ \bar{D}^*(2007)^0 \times B(D_{s0}(2317)^+ \rightarrow D_s^+ \pi^0))/\Gamma_{\text{total}}$				$\Gamma_{153}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.9 \pm 0.6 \pm 0.4</math></b>	<sup>1</sup> AUBERT,B	04s	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_{sJ}(2457)^+ \bar{D}^0)/\Gamma_{\text{total}}$				$\Gamma_{154}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>3.1 \pm 1.0</math> OUR AVERAGE</b>				
$4.3 \pm 1.6 \pm 1.3$	<sup>1</sup> AUBERT	06N	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$4.6 \pm 1.8 \pm 1.0$	<sup>2,3</sup> AUBERT,B	04s	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$2.1 \pm 1.1 \pm 0.5$	<sup>2,4</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses a missing-mass method in the events that one of the  $B$  mesons is fully reconstructed.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> AUBERT,B 04s reports  $[\Gamma(B^+ \rightarrow D_{sJ}(2457)^+ \bar{D}^0)/\Gamma_{\text{total}}] \times [B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0)] = (2.2 \pm 0.8 \pm 0.7 \pm 0.3) \times 10^{-3}$  which we divide by our best value  $B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0) = (48 \pm 11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> KROKOVNY 03B reports  $[\Gamma(B^+ \rightarrow D_{sJ}(2457)^+ \bar{D}^0)/\Gamma_{\text{total}}] \times [B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0)] = (1.0 \pm 0.5 \pm 0.1) \times 10^{-3}$  which we divide by our best value  $B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0) = (48 \pm 11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D_{sJ}(2457)^+ \bar{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \gamma))/\Gamma_{\text{total}}$				$\Gamma_{155}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.46 \pm 0.13</math> OUR AVERAGE</b>				
$0.48 \pm 0.19 \pm 0.04$	<sup>1,2</sup> AUBERT,B	04s	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$0.45 \pm 0.15 \pm 0.04$	<sup>1,3</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> AUBERT,B 04s reports  $(0.6 \pm 0.2 \pm 0.2 \pm 0.1) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow D_{sJ}(2457)^+ \bar{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \gamma))/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> KROKOVNY 03B reports  $(0.56 \pm 0.16 \pm 0.15 \pm 0.17) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow D_{sJ}(2457)^+ \bar{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \gamma))/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D_{sJ}(2457)^+ \bar{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \pi^+ \pi^-))/\Gamma_{\text{total}}$				$\Gamma_{156}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.22</b>	90	<sup>1</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_{sJ}(2457)^+ \bar{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \pi^0))/\Gamma_{\text{total}}$				$\Gamma_{157}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.27</b>	90	<sup>1</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_{sJ}(2457)^+ \bar{D}^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^{*+} \gamma))/\Gamma_{\text{total}}$				$\Gamma_{158}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.98</b>	90	<sup>1</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_{sJ}(2457)^+ \bar{D}^*(2007)^0)/\Gamma_{\text{total}}$				$\Gamma_{159}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>12.0 \pm 3.0</math> OUR AVERAGE</b>				
$11.2 \pm 2.6 \pm 2.0$	<sup>1</sup> AUBERT	06N	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$16 \pm 8 \pm 4$	<sup>2,3</sup> AUBERT,B	04s	BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses a missing-mass method in the events that one of the  $B$  mesons is fully reconstructed.

<sup>2</sup> AUBERT,B 04s reports  $[\Gamma(B^+ \rightarrow D_{sJ}(2457)^+ \bar{D}^*(2007)^0)/\Gamma_{\text{total}}] \times [B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0)] = (7.6 \pm 1.7 \pm 3.2 \pm 2.4) \times 10^{-3}$  which we divide by our best value  $B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0) = (48 \pm 11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D_{sJ}(2457)^+ \bar{D}^*(2007)^0 \times B(D_{sJ}(2457)^+ \rightarrow D_s^+ \gamma))/\Gamma_{\text{total}}$				$\Gamma_{160}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.4 \pm 0.4 \pm 0.6 \pm 0.4</math></b>	<sup>1</sup> AUBERT,B	04s	BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+))/\Gamma_{\text{total}}$				$\Gamma_{162}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.16 \pm 0.52 \pm 0.45</math></b>	<sup>1</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<2	90	AUBERT	03x	BABR Repl. by AUBERT 08B
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{D}^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+ + D^*(2010)^+ K^0))/\Gamma_{\text{total}}$				$\Gamma_{161}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>3.97 \pm 0.85 \pm 0.56</math></b>	<sup>1,2</sup> AUSHEV	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses $\Gamma(D^*(2007)^0 \rightarrow D^0 \pi^0) / \Gamma(D^*(2007)^0 \rightarrow D^0 \gamma) = 1.74 \pm 0.13$ and $\Gamma(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+) / \Gamma(D_{s1}(2536)^+ \rightarrow D^*(2010)^+ K^0) = 1.36 \pm 0.2$ .				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$$\Gamma(\bar{D}^*(2007)^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+))/\Gamma_{\text{total}} \quad \Gamma_{163}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.46 \pm 1.17 \pm 1.04</math></b>		<sup>1</sup> AUBERT	08B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<7	90	AUBERT	03X	BABR Repl. by AUBERT 08B
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^0 D_{s1}(2536)^+ \times B(D_{s1}(2536)^+ \rightarrow D^{*+} K^0))/\Gamma_{\text{total}} \quad \Gamma_{164}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.30 \pm 0.98 \pm 0.43</math></b>	<sup>1</sup> AUBERT	08B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^0 D_{sJ}(2700)^+ \times B(D_{sJ}(2700)^+ \rightarrow D^0 K^+))/\Gamma_{\text{total}} \quad \Gamma_{165}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.6 \pm 1.8</math> OUR AVERAGE</b>	Error includes scale factor of 1.7.		
$5.02 \pm 0.71 \pm 0.93$	<sup>1</sup> LEES	15c	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$11.3 \pm 2.2 \pm 1.4$	<sup>1</sup> BRODZICKA	08	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$2.8$			

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^* D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0)/\Gamma_{\text{total}} \quad \Gamma_{166}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.92 \pm 2.46 \pm 0.83</math></b>	<sup>1</sup> AUBERT	08B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^0 D_{sJ}(2573)^+, D_{sJ}^+ \rightarrow D^0 K^+)/\Gamma_{\text{total}} \quad \Gamma_{167}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.08 \pm 0.14 \pm 0.05</math></b>	<sup>1</sup> LEES	15c	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^* D_{sJ}(2573), D_{sJ}^+ \rightarrow D^0 K^+)/\Gamma_{\text{total}} \quad \Gamma_{168}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2</b>	90	AUBERT	03X	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

$$\Gamma(\bar{D}^*(2007)^0 D_{sJ}(2573), D_{sJ}^+ \rightarrow D^0 K^+)/\Gamma_{\text{total}} \quad \Gamma_{169}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5</b>	90	AUBERT	03X	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

$$\Gamma(\bar{D}^0 D_s^+)/\Gamma_{\text{total}} \quad \Gamma_{170}/\Gamma$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0076 \pm 0.0016</math> OUR AVERAGE</b>			
$0.0079 \pm 0.0017 \pm 0.0007$	<sup>1</sup> AUBERT	06N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.0068 \pm 0.0025 \pm 0.0006$	<sup>2</sup> GIBAUT	96	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$0.010 \pm 0.007 \pm 0.001$	<sup>3</sup> ALBRECHT	92G	ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 06N reports  $(0.77 \pm 0.15 \pm 0.13) \times 10^{-2}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.0462 \pm 0.0062$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> GIBAUT 96 reports  $0.0087 \pm 0.0027 \pm 0.0017$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.035$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> ALBRECHT 92G reports  $0.016 \pm 0.012 \pm 0.003$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^0$  branching ratios, e.g.,  $B(D^0 \rightarrow K^- \pi^+) = 3.71 \pm 0.25\%$ .

$$\Gamma(\bar{D}^*(2007)^0 D_s^+)/\Gamma_{\text{total}} \quad \Gamma_{171}/\Gamma$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0082 \pm 0.0017</math> OUR AVERAGE</b>			
$0.0078 \pm 0.0018 \pm 0.0007$	<sup>1</sup> AUBERT	06N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.011 \pm 0.004 \pm 0.001$	<sup>2</sup> GIBAUT	96	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$0.008 \pm 0.006 \pm 0.001$	<sup>3</sup> ALBRECHT	92G	ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 06N reports  $(0.76 \pm 0.15 \pm 0.13) \times 10^{-2}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.0462 \pm 0.0062$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> GIBAUT 96 reports  $0.0140 \pm 0.0043 \pm 0.0035$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.035$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> ALBRECHT 92G reports  $0.013 \pm 0.009 \pm 0.002$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 D_s^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ ,

which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^0$  and  $D^*(2007)^0$  branching ratios, e.g.,  $B(D^0 \rightarrow K^- \pi^+) = 3.71 \pm 0.25\%$  and  $B(D^*(2007)^0 \rightarrow D^0 \pi^0) = 55 \pm 6\%$ .

$$\Gamma(\bar{D}^*(2007)^0 D_s^{*+})/\Gamma_{\text{total}} \quad \Gamma_{172}/\Gamma$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0171 \pm 0.0024</math> OUR AVERAGE</b>			
$0.0167 \pm 0.0019 \pm 0.0015$	<sup>1</sup> AUBERT	06N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.024 \pm 0.009 \pm 0.002$	<sup>2</sup> GIBAUT	96	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$0.019 \pm 0.010 \pm 0.002$	<sup>3</sup> ALBRECHT	92G	ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 06N reports  $(1.62 \pm 0.22 \pm 0.18) \times 10^{-2}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.0462 \pm 0.0062$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> GIBAUT 96 reports  $0.0310 \pm 0.0088 \pm 0.0065$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.035$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> ALBRECHT 92G reports  $0.031 \pm 0.016 \pm 0.005$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{D}^*(2007)^0 D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^0$  and  $D^*(2007)^0$  branching ratios, e.g.,  $B(D^0 \rightarrow K^- \pi^+) = 3.71 \pm 0.25\%$  and  $B(D^*(2007)^0 \rightarrow D^0 \pi^0) = 55 \pm 6\%$ .

$$\Gamma(D_s^{*+} \bar{D}^{*0})/\Gamma_{\text{total}} \quad \Gamma_{173}/\Gamma$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>(2.73 \pm 0.93 \pm 0.68) \times 10^{-2}</math></b>	<sup>1</sup> AHMED	00B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AHMED 00B reports their experiment's uncertainties  $(\pm 0.78 \pm 0.48 \pm 0.68)\%$ , where the first error is statistical, the second is systematic, and the third is the uncertainty in the  $D_s \rightarrow \phi\pi$  branching fraction. We combine the first two in quadrature.

$$\Gamma(\bar{D}^*(2007)^0 D^*(2010)^+)/\Gamma_{\text{total}} \quad \Gamma_{174}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>8.1 \pm 1.2 \pm 1.2</math></b>		<sup>1</sup> AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<110	90	BARATE	98Q	ALEP $e^+ e^- \rightarrow Z$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$[\Gamma(\bar{D}^0 D^*(2010)^+) + \Gamma(\bar{D}^*(2007)^0 D^+)]/\Gamma_{\text{total}} \quad \Gamma_{175}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;130</b>	90	BARATE	98Q	ALEP $e^+ e^- \rightarrow Z$

$$\Gamma(\bar{D}^0 D^*(2010)^+)/\Gamma_{\text{total}} \quad \Gamma_{176}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.9 \pm 0.5</math> OUR AVERAGE</b>			
$3.6 \pm 0.5 \pm 0.4$	<sup>1</sup> AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$4.57 \pm 0.71 \pm 0.56$	<sup>1</sup> MAJUMDER	05	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^0 D^+)/\Gamma_{\text{total}} \quad \Gamma_{177}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>3.8 \pm 0.4</math> OUR AVERAGE</b>				
$3.85 \pm 0.31 \pm 0.38$		<sup>1</sup> ADACHI	08	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$3.8 \pm 0.6 \pm 0.5$		<sup>1</sup> AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.83 \pm 0.78 \pm 0.58$		<sup>1</sup> MAJUMDER	05	BELL Repl. by ADACHI 08
<67	90	BARATE	98Q	ALEP $e^+ e^- \rightarrow Z$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(\bar{D}^0 D^+ K^0)/\Gamma_{\text{total}} \quad \Gamma_{178}/\Gamma$$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.55 \pm 0.17 \pm 0.13</math></b>		<sup>1</sup> DEL-AMO-SA..11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.8	90	<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$$\Gamma(D^+ \bar{D}^*(2007)^0)/\Gamma_{\text{total}} \quad \Gamma_{179}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>6.3 \pm 1.4 \pm 1.0</math></b>	<sup>1</sup> AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

# Meson Particle Listings

$B^\pm$

$\Gamma(\bar{D}^*(2007)^0 D^+ K^0)/\Gamma_{\text{total}}$					$\Gamma_{180}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.06 \pm 0.31 \pm 0.30</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
<6.1	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^0 D^*(2010)^+ K^0)/\Gamma_{\text{total}}$					$\Gamma_{181}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>3.81 \pm 0.31 \pm 0.23</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$5.2^{+1.0}_{-0.9} \pm 0.7$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^*(2007)^0 D^*(2010)^+ K^0)/\Gamma_{\text{total}}$					$\Gamma_{182}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>9.17 \pm 0.83 \pm 0.90</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$7.8^{+2.3}_{-2.1} \pm 1.4$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^0 D^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{183}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>1.45 \pm 0.33</math> OUR AVERAGE</b>		Error includes scale factor of 2.6.			
$1.31 \pm 0.07 \pm 0.12$		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
$2.22 \pm 0.22 \pm 0.26_{-0.24}$		<sup>1</sup> BRODZICKA 08	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$1.17 \pm 0.21 \pm 0.15$		<sup>1</sup> CHISTOV 04	BELL	Repl. by BRODZICKA 08	
$1.9 \pm 0.3 \pm 0.3$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^*(2007)^0 D^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{184}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.26 \pm 0.16 \pm 0.17</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
<3.8	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^0 D^*(2007)^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{185}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>6.32 \pm 0.19 \pm 0.45</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$4.7 \pm 0.7 \pm 0.7$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^*(2007)^0 D^*(2007)^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{186}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>11.23 \pm 0.36 \pm 1.26</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$5.3^{+1.1}_{-1.0} \pm 1.2$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D^- D^+ K^+)/\Gamma_{\text{total}}$					$\Gamma_{187}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.22 \pm 0.05 \pm 0.05</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
<0.90	90	<sup>1</sup> CHISTOV 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
<0.4	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D^- D^*(2010)^+ K^+)/\Gamma_{\text{total}}$					$\Gamma_{188}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.63 \pm 0.09 \pm 0.06</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
<0.7	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D^*(2010)^- D^+ K^+)/\Gamma_{\text{total}}$					$\Gamma_{189}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>0.60 \pm 0.10 \pm 0.08</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$1.5 \pm 0.3 \pm 0.2$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D^*(2010)^- D^*(2010)^+ K^+)/\Gamma_{\text{total}}$					$\Gamma_{190}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.32 \pm 0.13 \pm 0.12</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
<1.8	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma((\bar{D}+\bar{D}^*)(D+D^*)K)/\Gamma_{\text{total}}$					$\Gamma_{191}/\Gamma$
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>4.05 \pm 0.11 \pm 0.28</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
$3.5 \pm 0.3 \pm 0.5$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D_s^+ \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{192}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.6^{+0.6}_{-0.5} \pm 0.1</math></b>		<sup>1</sup> AUBERT	07M	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.	• • •				
<16	90	<sup>2</sup> ALEXANDER 93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> AUBERT 07M reports $[\Gamma(B^+ \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)] = (7.0^{+2.4+0.6}_{-2.1-0.8}) \times 10^{-7}$ which we divide by our best value $B(D_s^+ \rightarrow \phi \pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> ALEXANDER 93B reports $< 2.0 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$[\Gamma(D_s^+ \pi^0) + \Gamma(D_s^+ \pi^0)]/\Gamma_{\text{total}}$					$(\Gamma_{192} + \Gamma_{193})/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;5 \times 10^{-4}</math></b>	90	<sup>1</sup> ALBRECHT 93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALBRECHT 93E reports $< 0.9 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \pi^0) + \Gamma(B^+ \rightarrow D_s^+ \pi^0)]/\Gamma_{\text{total}} \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^{*+} \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{193}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2.6 \times 10^{-4}</math></b>	90	<sup>1</sup> ALEXANDER 93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALEXANDER 93B reports $< 3.2 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi \pi^+)]$ assuming $B(D_s^{*+} \rightarrow \phi \pi^+) = 0.037$ , which we rescale to our best value $B(D_s^{*+} \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^+ \eta)/\Gamma_{\text{total}}$					$\Gamma_{194}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;4 \times 10^{-4}</math></b>	90	<sup>1</sup> ALEXANDER 93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALEXANDER 93B reports $< 4.6 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \eta)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^{*+} \eta)/\Gamma_{\text{total}}$					$\Gamma_{195}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6 \times 10^{-4}</math></b>	90	<sup>1</sup> ALEXANDER 93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALEXANDER 93B reports $< 7.5 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \eta)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi \pi^+)]$ assuming $B(D_s^{*+} \rightarrow \phi \pi^+) = 0.037$ , which we rescale to our best value $B(D_s^{*+} \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^+ \rho^0)/\Gamma_{\text{total}}$					$\Gamma_{196}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;3.0 \times 10^{-4}</math></b>	90	<sup>1</sup> ALEXANDER 93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALEXANDER 93B reports $< 3.7 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \rho^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					



$\Gamma(D_s^+ \rho^0) + \Gamma(D_s^+ \bar{K}^*(892)^0)/\Gamma_{\text{total}}$  ( $\Gamma_{196} + \Gamma_{206}$ )/ $\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.0 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93E reports $< 3.4 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \rho^0) + \Gamma(B^+ \rightarrow D_s^+ \bar{K}^*(892)^0)]/\Gamma_{\text{total}} \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^{*+} \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{197}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 4.8 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \rho^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $[\Gamma(D_s^{*+} \rho^0) + \Gamma(D_s^{*+} \bar{K}^*(892)^0)]/\Gamma_{\text{total}}$  ( $\Gamma_{197} + \Gamma_{208}$ )/ $\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.2 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93E reports $< 2.0 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \rho^0) + \Gamma(B^+ \rightarrow D_s^{*+} \bar{K}^*(892)^0)]/\Gamma_{\text{total}} \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^+ \omega)/\Gamma_{\text{total}}$   $\Gamma_{198}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.0 \times 10^{-3}$	90	<sup>2</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 4.8 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \omega)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				
<sup>2</sup> ALBRECHT 93E reports $< 3.4 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \omega)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^{*+} \omega)/\Gamma_{\text{total}}$   $\Gamma_{199}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<6 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.1 \times 10^{-3}$	90	<sup>2</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 6.8 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \omega)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				
<sup>2</sup> ALBRECHT 93E reports $< 1.9 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \omega)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^+ a_1(1260)^0)/\Gamma_{\text{total}}$   $\Gamma_{200}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93E reports $< 3.0 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ a_1(1260)^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^{*+} a_1(1260)^0)/\Gamma_{\text{total}}$   $\Gamma_{201}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.3 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93E reports $< 2.2 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} a_1(1260)^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^+ \phi)/\Gamma_{\text{total}}$   $\Gamma_{202}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$1.7^{+1.1}_{-0.7} \pm 0.2$		<sup>1</sup> AAIJ 13R	13R LHCB	$pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.9$	90	<sup>2</sup> AUBERT 06F	06F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$<1000$	90	<sup>3</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$< 260$	90	<sup>4</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 13R reports  $(1.87^{+1.25}_{-0.73} \pm 0.19 \pm 0.32) \times 10^{-6}$  from a measurement of  $[\Gamma(B^+ \rightarrow D_s^+ \phi)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^0 D_s^+)]$  assuming  $B(B^+ \rightarrow \bar{D}^0 D_s^+) = (10.0 \pm 1.7) \times 10^{-3}$ , which we rescale to our best value  $B(B^+ \rightarrow \bar{D}^0 D_s^+) = (9.0 \pm 0.9) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> ALBRECHT 93E reports  $< 1.7 \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow D_s^+ \phi)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

<sup>4</sup> ALEXANDER 93B reports  $< 3.1 \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow D_s^+ \phi)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

 $\Gamma(D_s^{*+} \phi)/\Gamma_{\text{total}}$   $\Gamma_{203}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.2 \times 10^{-5}$	90	<sup>1</sup> AUBERT 06F	06F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.3 \times 10^{-3}$	90	<sup>2</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$<3.5 \times 10^{-4}$	90	<sup>3</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> ALBRECHT 93E reports $< 2.1 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \phi)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				
<sup>3</sup> ALEXANDER 93B reports $< 4.2 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \phi)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^+ \bar{K}^0)/\Gamma_{\text{total}}$   $\Gamma_{204}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.5 \times 10^{-3}$	90	<sup>2</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 10.3 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \bar{K}^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				
<sup>2</sup> ALBRECHT 93E reports $< 2.5 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \bar{K}^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^{*+} \bar{K}^0)/\Gamma_{\text{total}}$   $\Gamma_{205}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<9 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.9 \times 10^{-3}$	90	<sup>2</sup> ALBRECHT 93E	93E ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 10.9 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \bar{K}^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				
<sup>2</sup> ALBRECHT 93E reports $< 3.1 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \bar{K}^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^+ \bar{K}^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{206}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.4 \times 10^{-6}$	90	AAIJ 13R	LHCB	$pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 4.4 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^+ \bar{K}^*(892)^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

 $\Gamma(D_s^+ K^{*0})/\Gamma_{\text{total}}$   $\Gamma_{207}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$<3.5$	90	AAIJ 13R	LHCB	$pp$ at 7 TeV

 $\Gamma(D_s^{*+} \bar{K}^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{208}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.5 \times 10^{-4}$	90	<sup>1</sup> ALEXANDER 93B	93B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALEXANDER 93B reports $< 4.3 \times 10^{-4}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*+} \bar{K}^*(892)^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$ assuming $B(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .				

# Meson Particle Listings

$B^\pm$

$(D_s^- \pi^+ K^+)/\Gamma_{\text{total}}$				$\Gamma_{209}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>1.80±0.22 OUR AVERAGE</b>					
$1.71^{+0.08}_{-0.07} \pm 0.25$		<sup>1</sup> WIECHCZYN...09	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
$2.02 \pm 0.13 \pm 0.38$		<sup>1</sup> AUBERT 08G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<7	90	<sup>2</sup> ALBRECHT 93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> ALBRECHT 93E reports $< 1.1 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^- \pi^+ K^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^{*-} \pi^+ K^+)/\Gamma_{\text{total}}$			$\Gamma_{210}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.45 ± 0.24 OUR AVERAGE</b>				
$1.31^{+0.13}_{-0.12} \pm 0.28$		<sup>1</sup> WIECHCZYN...09	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.67 \pm 0.16 \pm 0.35$		<sup>1</sup> AUBERT 08G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<10	90	<sup>2</sup> ALBRECHT 93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> ALBRECHT 93E reports $< 1.6 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*-} \pi^+ K^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .				

$\Gamma(D_s^- \pi^+ K^*(892)^+)/\Gamma_{\text{total}}$				$\Gamma_{211}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;5 × 10<sup>-3</sup></b>	90	<sup>1</sup> ALBRECHT 93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALBRECHT 93E reports $< 8.6 \times 10^{-3}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^- \pi^+ K^*(892)^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^{*-} \pi^+ K^*(892)^+)/\Gamma_{\text{total}}$				$\Gamma_{212}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93E reports $< 1.1 \times 10^{-2}$ from a measurement of $[\Gamma(B^+ \rightarrow D_s^{*-} \pi^+ K^*(892)^+)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$ assuming $B(D_s^+ \rightarrow \phi \pi^+) = 0.027$ , which we rescale to our best value $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .				

$\Gamma(D_s^- K^+ K^+)/\Gamma_{\text{total}}$				$\Gamma_{213}/\Gamma$	
VALUE (units $10^{-5}$ )		DOCUMENT ID	TECN	COMMENT	
<b>0.97±0.21 OUR AVERAGE</b>					
$0.93 \pm 0.22 \pm 0.10$		<sup>1</sup> WIECHCZYN...15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
$1.1 \pm 0.4 \pm 0.2$		<sup>1</sup> AUBERT 08G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D_s^- K^+ K^+)/\Gamma(D_s^- \pi^+ K^+)$				$\Gamma_{213}/\Gamma_{209}$
VALUE		DOCUMENT ID	TECN	COMMENT
<b>0.054±0.013±0.006</b>		WIECHCZYN...15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(D_s^{*-} K^+ K^+)/\Gamma_{\text{total}}$				$\Gamma_{214}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.15</b>	90	<sup>1</sup> AUBERT 08G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\langle \eta_c K^+ \rangle / \Gamma_{\text{total}}$					$\Gamma_{215} / \Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b>0.96<math>\pm</math>0.11 OUR AVERAGE</b>					
0.87 $\pm$ 0.15		<sup>1,2</sup> AUBERT	06E	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
1.20 $^{+0.24}_{-0.19}$ $\pm$ 0.13		<sup>3</sup> AUBERT,B	05L	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
1.25 $\pm$ 0.14 $^{+0.39}_{-0.40}$		<sup>4</sup> FANG	03	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
0.69 $^{+0.26}_{-0.21}$ $\pm$ 0.22		<sup>5</sup> EDWARDS	01	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.02 $\pm$ 0.12 $\pm$ 0.07		<sup>2,6</sup> AUBERT,B	04B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Perform measurements of absolute branching fractions using a missing mass technique.					
<sup>2</sup> The ratio of $B(K^\pm \rightarrow K^\pm \eta_c) B(\eta_c \rightarrow K \bar{K} \pi) = (7.4 \pm 0.5 \pm 0.7) \times 10^{-5}$ reported in AUBERT,B 04B and $B(K^\pm \rightarrow K^\pm \eta_c) = (8.7 \pm 1.5) \times 10^{-3}$ reported in AUBERT 06E contribute to the determination of $B(\eta_c \rightarrow K \bar{K} \pi)$ , which is used by others for normalization.					
<sup>3</sup> AUBERT,B 05L reports $[\Gamma(B^+ \rightarrow \eta_c K^+) / \Gamma_{\text{total}}] \times [B(\eta_c(1S) \rightarrow p \bar{p})] = (1.8^{+0.3}_{-0.2} \pm 0.2) \times 10^{-6}$ which we divide by our best value $B(\eta_c(1S) \rightarrow p \bar{p}) = (1.50 \pm 0.16) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>4</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>5</sup> EDWARDS 01 assumes equal production of $B^0$ and $B^+$ at the $\Upsilon(4S)$ . The correlated uncertainties (28.3)% from $B(J/\psi(1S) \rightarrow \gamma \eta_c)$ in those modes have been accounted for.					
<sup>6</sup> AUBERT,B 04B reports $[\Gamma(B^+ \rightarrow \eta_c K^+) / \Gamma_{\text{total}}] \times [B(\eta_c(1S) \rightarrow K \bar{K} \pi)] = (0.074 \pm 0.005 \pm 0.007) \times 10^{-3}$ which we divide by our best value $B(\eta_c(1S) \rightarrow K \bar{K} \pi) = (7.3 \pm 0.5) \times 10^{-2}$ .					

<sup>5</sup> EDWARDS 01 assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ . The correlated uncertainties (28.3)% from  $B(J/\psi(1S) \rightarrow \gamma \eta_c)$  in those modes have been accounted for.

<sup>6</sup> AUBERT,B 04B reports  $[\Gamma(B^+ \rightarrow \eta_c K^+)/\Gamma_{\text{total}}] \times [B(\eta_c(1S) \rightarrow K \bar{K} \pi)] = (0.074 \pm 0.005 \pm 0.007) \times 10^{-3}$  which we divide by our best value  $B(\eta_c(1S) \rightarrow K \bar{K} \pi) = (7.3 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(B^+ \rightarrow \eta_c K^+)/\Gamma_{\text{total}} \times \Gamma(\eta_c(1S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$				$\Gamma_{215}/\Gamma \times \Gamma_{47}^{\eta_c(1S)}/\Gamma_{\eta_c(1S)}$	
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
$0.22 \pm 0.09 \pm 0.04$ $-0.07 - 0.02$		<sup>1</sup> WICHT	08	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\eta_c K^+, \eta_c \rightarrow K_S^0 K^\mp \pi^\pm)/\Gamma_{\text{total}}$				$\Gamma_{216}/\Gamma$	
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b>26.7<math>\pm</math>1.4<math>^{+5.7}_{-5.5}</math></b>	1,2	VINOKUROVA 11	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^0$ and $B^+$ from Upsilon(4S) decays.					
<sup>2</sup> VINOKUROVA 11 reports $(26.7 \pm 1.4^{+2.9}_{-2.6} \pm 4.9) \times 10^{-6}$ , where the first uncertainty is statistical, the second is due to systematics, and the third comes from interference of $\eta_c(1S) \rightarrow K_S^0 K^\pm \pi^\mp$ with nonresonant $K_S^0 K^\pm \pi^\mp$ . We combined both systematic uncertainties to single values.					

$\Gamma(\eta_c K^*(892^+)/\Gamma_{\text{total}}$				$\Gamma_{217}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
$1.0^{+0.5}_{-0.4} \pm 0.1$	1,2 AUBERT	07AV BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> AUBERT 07AV reports $[\Gamma(B^+ \rightarrow \eta_c K^*(892^+)/\Gamma_{\text{total}}] \times [B(\eta_c(1S) \rightarrow p\bar{p})] = (1.57^{+0.56+0.45}_{-0.46-0.36}) \times 10^{-6}$ which we divide by our best value $B(\eta_c(1S) \rightarrow p\bar{p}) = (1.50 \pm 0.16) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\eta_c K^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{218}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.9 \times 10^{-4}$	90	VINOKUROVA 15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

$\Gamma(\eta_c K^+ \omega(782))/\Gamma_{\text{total}}$				$\Gamma_{219}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<5.3 \times 10^{-4}$	90	VINOKUROVA 15	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(\eta_c K^+ \eta)/\Gamma_{\text{total}}$				$\Gamma_{220}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.2 \times 10^{-4}$	90	VINOKUROVA 15	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(\eta_C K^+ \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{221}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6.2 \times 10^{-5}$	90	VINOKUROVA 15	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(\eta_c(2S) K^+)/\Gamma_{\text{total}}$				$\Gamma_{222}/\Gamma$	
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>3.4 \pm 1.8 \pm 0.3</math></b>	<sup>1</sup>	AUBERT	06E BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Perform measurements of absolute branching fractions using a missing mass technique.					

$\Gamma(\eta_c(2S) K^+, \eta_c \rightarrow p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{223}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.06 \times 10^{-7}</math></b>	95	<sup>1</sup> AAIJ 13s	LHCB	$p p$ at 7 TeV	
<sup>1</sup> Measured relative to $B^+ \rightarrow J/\psi K^+$ decay with charmonia reconstructed in $p \bar{p}$ final state and using $B(B^+ \rightarrow J/\psi K^+) = (1.013 \pm 0.034) \times 10^{-3}$ and $B(J/\psi \rightarrow p \bar{p}) = (2.17 \pm 0.07) \times 10^{-3}$ .					

$\Gamma(B^+ \rightarrow h_c(1P) K^+)/\Gamma_{\text{total}} \times \Gamma(h_c(1P) \rightarrow \eta_c(1S) \gamma)/\Gamma_{\text{total}}$			$\Gamma_{303}/\Gamma \times \Gamma_{h_c(1P)}^{h_c(1P)}/\Gamma_{h_c(1P)}$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.48</b>	90	<sup>1</sup> AUBERT 08AB	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses the production ratio of $(B^+B^-)/(B^0\bar{B}^0) = 1.026 \pm 0.032$ at $\Upsilon(4S)$ .				

$\Gamma(B^+ \rightarrow \eta_c(2S) K^+)/\Gamma_{\text{total}} \times \Gamma(\eta_c(2S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$			$\Gamma_{222}/\Gamma \times \Gamma_{15}^{\eta_c(2S)}/\Gamma_{\eta_c(2S)}$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.18</b>	90	<sup>1</sup> WICHT 08	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\eta_c(2S) K^+, \eta_c \rightarrow K_S^0 K^\mp \pi^\pm)/\Gamma_{\text{total}}$				$\Gamma_{224}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
$3.4^{+2.2+0.5}_{-1.5-0.4}$	1,2	VINOKUROVA 11	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^0$ and $B^+$ from Upsilon(4S) decays.				
<sup>2</sup> The first uncertainty includes both statistical and interference effects while the second is due to systematics.				

$\Gamma(J/\psi(1S)K^+)/\Gamma_{\text{total}}$		$\Gamma_{255}/\Gamma$		
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.26 ± 0.31 OUR FIT</b>				
<b>10.24 ± 0.35 OUR AVERAGE</b>				
8.1 ± 1.3 ± 0.7		<sup>1</sup> AUBERT	06E	BABR $e^+e^- \rightarrow \Upsilon(4S)$
10.61 ± 0.15 ± 0.48		<sup>2</sup> AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
10.4 ± 1.1 ± 0.1		<sup>3</sup> AUBERT,B	05L	BABR $e^+e^- \rightarrow \Upsilon(4S)$
10.1 ± 0.2 ± 0.7		<sup>2</sup> ABE	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
10.2 ± 0.8 ± 0.7		<sup>2</sup> JESSOP	97	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
9.24 ± 3.04 ± 0.05		<sup>4</sup> BORTOLETTO	92	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
8.09 ± 3.50 ± 0.04	6	<sup>5</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- 10.1 ± 0.3 ± 0.5     <sup>2</sup> AUBERT     02     BABR     Repl. by AUBERT 05J
- 11.0 ± 1.5 ± 0.9     59     <sup>2</sup> ALAM     94     CLE2     Repl. by JESSOP 97
- 22 ± 10 ± 2     <sup>6</sup> BUSKULIC     92G     ALEP      $e^+e^- \rightarrow Z$
- 7 ± 4     3     <sup>6</sup> ALBRECHT     87D     ARG      $e^+e^- \rightarrow \Upsilon(4S)$
- 10 ± 7 ± 2     3     <sup>7</sup> BEBEK     87     CLEO      $e^+e^- \rightarrow \Upsilon(4S)$
- 9 ± 5     3     <sup>8</sup> ALAM     86     CLEO      $e^+e^- \rightarrow \Upsilon(4S)$
- <sup>1</sup> Perform measurements of absolute branching fractions using a missing mass technique.
- <sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>3</sup> AUBERT,B 05L reports  $[\Gamma(B^+ \rightarrow J/\psi(1S)K^+)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow p\bar{p})] = (2.2 \pm 0.2 \pm 0.1) \times 10^{-6}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow p\bar{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>4</sup> BORTOLETTO 92 reports  $(8 \pm 2 \pm 2) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S)K^+)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>5</sup> ALBRECHT 90J reports  $(7 \pm 3 \pm 1) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S)K^+)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>6</sup> ALBRECHT 87D assume  $B^+B^-/B^0\bar{B}^0$  ratio is 55/45. Superseded by ALBRECHT 90J.
- <sup>7</sup> BEBEK 87 value has been updated in BERKELMAN 91 to use same assumptions as noted for BORTOLETTO 92.
- <sup>8</sup> ALAM 86 assumes  $B^\pm/B^0$  ratio is 60/40.

$\Gamma(\eta_c K^+)/\Gamma(J/\psi(1S)K^+)$		$\Gamma_{215}/\Gamma_{255}$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.84 ± 0.10 OUR AVERAGE</b>				
0.82 ± 0.06 ± 0.09		<sup>1</sup> AAIJ	13s	LHCB $pp$ at 7 TeV
1.33 ± 0.10 ± 0.43		<sup>2</sup> AUBERT,B	04B	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 13s reports  $[\Gamma(B^+ \rightarrow \eta_c K^+)/\Gamma(B^+ \rightarrow J/\psi(1S)K^+)] \times [B(\eta_c(1S) \rightarrow p\bar{p})] / [B(J/\psi(1S) \rightarrow p\bar{p})] = 0.578 \pm 0.035 \pm 0.026$  which we multiply or divide by our best values  $B(\eta_c(1S) \rightarrow p\bar{p}) = (1.50 \pm 0.16) \times 10^{-3}$ ,  $B(J/\psi(1S) \rightarrow p\bar{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> Uses BABAR measurement of  $B(B^+ \rightarrow J/\psi K^+) = (10.1 \pm 0.3 \pm 0.5) \times 10^{-4}$ .

$\Gamma(B^+ \rightarrow J/\psi(1S)K^+)/\Gamma_{\text{total}} \times \Gamma(J/\psi(1S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$		$\Gamma_{255}/\Gamma \times \Gamma_{222}^{(1S)}/\Gamma_{J/\psi(1S)}$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.16</b>				
	90	<sup>1</sup> WICHT	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(J/\psi(1S)K^+\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{257}/\Gamma$		
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.81 ± 0.13 OUR AVERAGE</b>				
Error includes scale factor of 2.5. See the ideogram below.				
0.716 ± 0.010 ± 0.060		<sup>1</sup> GULER	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
1.16 ± 0.07 ± 0.09		<sup>2</sup> AUBERT	05R	BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.69 ± 0.18 ± 0.12		<sup>2</sup> ACOSTA	02F	CDF $p\bar{p}$ 1.8 TeV
1.39 ± 0.81 ± 0.01		<sup>3</sup> BORTOLETTO	92	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
1.39 ± 0.91 ± 0.01	6	<sup>4</sup> ALBRECHT	87D	ARG $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

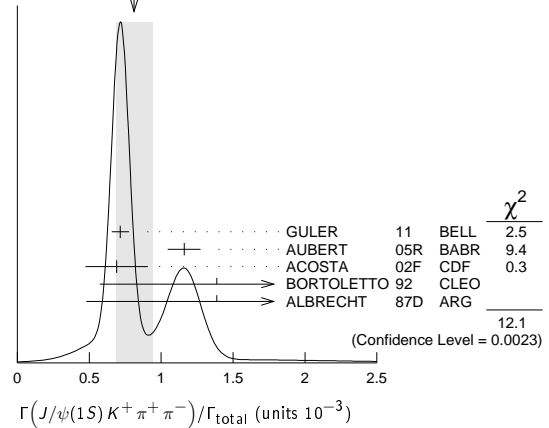
<1.8     90     <sup>5</sup> ALBRECHT     90J     ARG      $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>2</sup> ACOSTA 02F uses as reference of  $B(B \rightarrow J/\psi(1S)K^+) = (10.1 \pm 0.6) \times 10^{-4}$ . The second error includes the systematic error and the uncertainties of the branching ratio.
- <sup>3</sup> BORTOLETTO 92 reports  $(1.2 \pm 0.6 \pm 0.4) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S)K^+\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>4</sup> ALBRECHT 87D reports  $(1.2 \pm 0.8) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S)K^+\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is

the systematic error from using our best value. They actually report  $0.0011 \pm 0.0007$  assuming  $B^+B^-/B^0\bar{B}^0$  ratio is 55/45. We rescale to 50/50. Analysis explicitly removes  $B^+ \rightarrow \psi(2S)K^+$ .

- <sup>5</sup> ALBRECHT 90J reports  $< 1.6 \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S)K^+\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = 5.971 \times 10^{-2}$ . Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

WEIGHTED AVERAGE  
0.81±0.13 (Error scaled by 2.5)



$\Gamma(J/\psi(1S)K^+K^-K^+)/\Gamma_{\text{total}}$		$\Gamma_{258}/\Gamma$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>33.7 ± 2.5 ± 1.4</b>				
		LEES	15	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(h_c(1P)K^+, h_c \rightarrow J/\psi\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{225}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.4 × 10<sup>-6</sup></b>				
	90	<sup>1</sup> AUBERT	05R	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3730)^0K^+, X^0 \rightarrow \eta_c\eta)/\Gamma_{\text{total}}$		$\Gamma_{226}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.6 × 10<sup>-5</sup></b>				
	90	VINOKUROVA	15	BELL $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3730)^0K^+, X^0 \rightarrow \eta_c\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{227}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.7 × 10<sup>-6</sup></b>				
	90	VINOKUROVA	15	BELL $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3872)K^+)/\Gamma_{\text{total}}$		$\Gamma_{228}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.2 × 10<sup>-4</sup></b>				
	90	<sup>1</sup> AUBERT	06E	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Perform measurements of absolute branching fractions using a missing mass technique.

$\Gamma(B^+ \rightarrow X(3872)K^+)/\Gamma_{\text{total}} \times \Gamma(X(3872) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$		$\Gamma_{228}/\Gamma \times \Gamma_{\gamma}^{X(3872)}/\Gamma_{X(3872)}$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.24</b>				
	90	<sup>1</sup> WICHT	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872)K^+, X \rightarrow J/\psi\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{230}/\Gamma$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>8.6 ± 0.8 OUR AVERAGE</b>				
Error includes scale factor of 1.1.				
8.63 ± 0.82 ± 0.52		<sup>1</sup> CHOI	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
8.4 ± 1.5 ± 0.7		<sup>1</sup> AUBERT	08Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

10.1 ± 2.5 ± 1.0     <sup>1</sup> AUBERT     06     BABR     Repl. by AUBERT 08Y

12.8 ± 4.1     <sup>1</sup> AUBERT     05R     BABR     Repl. by AUBERT 06

12.5 ± 2.8 ± 0.5     <sup>2</sup> CHOI     03     BELL     Repl. by CHOI 11

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>2</sup> CHOI 03 reports  $[\Gamma(B^+ \rightarrow X(3872)K^+, X \rightarrow J/\psi\pi^+\pi^-)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow \psi(2S)K^+)] = 0.0200 \pm 0.0038 \pm 0.0023$  which we multiply by our best value  $B(B^+ \rightarrow \psi(2S)K^+) = (6.26 \pm 0.24) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(X(3872)K^+, X \rightarrow J/\psi\gamma)/\Gamma_{\text{total}}$		$\Gamma_{231}/\Gamma$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.1 ± 0.4 OUR AVERAGE</b>				
Error includes scale factor of 1.1.				
1.78 ± 0.48 ± 0.12		<sup>1</sup> BHARDWAJ	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
2.8 ± 0.8 ± 0.1		<sup>2</sup> AUBERT	09B	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.3 ± 1.0 ± 0.3     <sup>1</sup> AUBERT,BE     06M     BABR     Repl. by AUBERT 09B

# Meson Particle Listings

$B^\pm$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872) K^*(892)^+, X \rightarrow J/\psi \gamma)/\Gamma_{\text{total}}$   $\Gamma_{246}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.8</b>	90	<sup>1</sup> AUBERT	09B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872) K^+, X \rightarrow \psi(2S) \gamma)/\Gamma_{\text{total}}$   $\Gamma_{232}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>4 ±4 OUR AVERAGE</b>	Error includes scale factor of 2.5.		
$0.83^{+1.98}_{-1.83} \pm 0.44$	<sup>1,2</sup> BHARDWAJ	11 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$9.5 \pm 2.7 \pm 0.6$	<sup>3</sup> AUBERT	09B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> BHARDWAJ 11 measurement is equivalent to a limit of  $< 3.45 \times 10^{-6}$  at 90% CL.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872) K^*(892)^+, X \rightarrow \psi(2S) \gamma)/\Gamma_{\text{total}}$   $\Gamma_{247}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;28</b>	90	<sup>1</sup> AUBERT	09B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872) K^+, X \rightarrow D^0 \bar{D}^0)/\Gamma_{\text{total}}$   $\Gamma_{234}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.0 × 10<sup>-5</sup></b>	90	<sup>1</sup> CHISTOV	04 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872) K^+, X \rightarrow D^+ D^-)/\Gamma_{\text{total}}$   $\Gamma_{235}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.0 × 10<sup>-5</sup></b>	90	<sup>1</sup> CHISTOV	04 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872) K^+, X \rightarrow D^0 \bar{D}^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{236}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.02 \pm 0.31^{+0.21}_{-0.29}</math></b>		<sup>1</sup> GOKHROO	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<0.6 90 <sup>2</sup> CHISTOV 04 BELL Repl. by GOKHROO 06  
<sup>1</sup> Measure the near-threshold enhancements in the  $(D^0 \bar{D}^0 \pi^0)$  system at a mass  $3875.2 \pm 0.7^{+0.3}_{-1.6} \pm 0.8$  MeV/c<sup>2</sup>.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872) K^+, X \rightarrow \bar{D}^{*0} D^0)/\Gamma_{\text{total}}$   $\Gamma_{237}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.85 \pm 0.26</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.		
$0.77 \pm 0.16 \pm 0.10$	<sup>1</sup> AUSHEV	10 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.67 \pm 0.36 \pm 0.47$	<sup>1</sup> AUBERT	08B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{238}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.0 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3872)^0 K^+, X^0 \rightarrow \eta_c \omega(782))/\Gamma_{\text{total}}$   $\Gamma_{239}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.9 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3915)^0 K^+, X^0 \rightarrow \eta_c \eta)/\Gamma_{\text{total}}$   $\Gamma_{240}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.3 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3915)^0 K^+, X^0 \rightarrow \eta_c \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{241}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.8 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(4014)^0 K^+, X^0 \rightarrow \eta_c \eta)/\Gamma_{\text{total}}$   $\Gamma_{242}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.9 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(4014)^0 K^+, X^0 \rightarrow \eta_c \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{243}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.2 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3900)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{244}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.7 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(4020)^0 K^+, X^0 \rightarrow \eta_c \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{245}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.6 × 10<sup>-5</sup></b>	90	VINOKUROVA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(3872) K^+, X \rightarrow J/\psi(1S) \eta)/\Gamma_{\text{total}}$   $\Gamma_{233}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.7 × 10<sup>-6</sup></b>	90	<sup>1</sup> AUBERT	04Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872)^+ K^0, X^+ \rightarrow J/\psi(1S) \pi^+ \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{248}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 6.1</b>	90	<sup>1,2</sup> CHOI	11 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
<22	90	<sup>3</sup> AUBERT	05B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes  $\pi^+ \pi^0$  originates from  $\rho^+$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The isovector- $X$  hypothesis is excluded with a likelihood test at  $1 \times 10^{-4}$  level.

$\Gamma(X(3872) K^0 \pi^+, X \rightarrow J/\psi(1S) \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{249}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>10.6 \pm 3.0 \pm 0.9</math></b>	BALA	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(X(4430)^+ K^0, X^+ \rightarrow J/\psi \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{250}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.5</b>	95	<sup>1</sup> AUBERT	09AA BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(4430)^+ K^0, X^+ \rightarrow \psi(2S) \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{251}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.7</b>	95	<sup>1</sup> AUBERT	09AA BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(4260)^0 K^+, X^0 \rightarrow J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{252}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;29</b>	95	<sup>1</sup> AUBERT	06 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3915) K^+, X \rightarrow J/\psi \gamma)/\Gamma_{\text{total}}$   $\Gamma_{253}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;14</b>	90	<sup>1</sup> AUBERT,BE	06M BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3930)^0 K^+, X^0 \rightarrow J/\psi \gamma)/\Gamma_{\text{total}}$   $\Gamma_{254}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5</b>	90	<sup>1</sup> AUBERT,BE	06M BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(J/\psi(1S) K^0 \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{256}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
$1.101 \pm 0.021$	<sup>1</sup> AUBERT	09AA BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Does not report systematic uncertainties.

$\Gamma(J/\psi(1S) K^*(892)^+)/\Gamma_{\text{total}}$   $\Gamma_{260}/\Gamma$

For polarization information see the Listings at the end of the “ $B^0$  Branching Ratios” section.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.43 \pm 0.08</math> OUR FIT</b>				
<b><math>1.43 \pm 0.08</math> OUR AVERAGE</b>				

$1.78^{+0.36}_{-0.32} \pm 0.02$	<sup>1,2</sup> AUBERT	07av BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.454 \pm 0.047 \pm 0.097$	<sup>2</sup> AUBERT	05J BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.28 \pm 0.07 \pm 0.14$	<sup>2</sup> ABE	02N BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.41 \pm 0.23 \pm 0.24$	<sup>2</sup> JESSOP	97 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.58 \pm 0.47 \pm 0.27$	<sup>3</sup> ABE	96H CDF	$p\bar{p}$ at 1.8 TeV
$1.50 \pm 1.08 \pm 0.01$	<sup>4</sup> BORTOLETTO	92 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.85 \pm 1.30 \pm 0.01$	<sup>5</sup> ALBRECHT	90J ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
$1.37 \pm 0.09 \pm 0.11$	<sup>2</sup> AUBERT	02 BABR	Repl. by AUBERT 05J
$1.78 \pm 0.51 \pm 0.23$	<sup>2</sup> ALAM	94 CLE2	Sup. by JESSOP 97

<sup>1</sup> AUBERT 07av reports  $[\Gamma(B^+ \rightarrow J/\psi(1S) K^*(892)^+)/\Gamma_{\text{total}}] \times [\text{B}(J/\psi(1S) \rightarrow p\bar{p})] = (3.78^{+0.72+0.28}_{-0.64-0.23}) \times 10^{-6}$  which we divide by our best value  $\text{B}(J/\psi(1S) \rightarrow p\bar{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment’s error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> ABE 96H assumes that  $\text{B}(B^+ \rightarrow J/\psi K^+) = (1.02 \pm 0.14) \times 10^{-3}$ .  
<sup>4</sup> BORTOLETTO 92 reports  $(1.3 \pm 0.9 \pm 0.3) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S) K^*(892)^+)/\Gamma_{\text{total}}] \times [\text{B}(J/\psi(1S) \rightarrow e^+ e^-)]$  assuming  $\text{B}(J/\psi(1S) \rightarrow e^+ e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $\text{B}(J/\psi(1S) \rightarrow e^+ e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment’s error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>5</sup> ALBRECHT 90J reports  $(1.6 \pm 1.1 \pm 0.3) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow J/\psi(1S) K^*(892)^+)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+ e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+ e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+ e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) K^*(892)^+)/\Gamma(J/\psi(1S) K^+)$   $\Gamma_{260}/\Gamma_{255}$**

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.39±0.09 OUR AVERAGE</b>			
1.37±0.05±0.08	AUBERT	05J	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
1.45±0.20±0.17	<sup>1</sup> JESSOP	97	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
1.92±0.60±0.17	ABE	96Q	CDF $p\bar{p}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.37±0.10±0.08 <sup>2</sup> AUBERT 02 BABR Repl. by AUBERT 05J

<sup>1</sup> JESSOP 97 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The measurement is actually measured as an average over kaon charged and neutral states.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) K(1270)^+)/\Gamma_{\text{total}}$   $\Gamma_{261}/\Gamma$**

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.80±0.34±0.39</b>	<sup>1</sup> ABE	01L	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the PDG value of  $B(B^+ \rightarrow J/\psi(1S) K^+) = (1.00 \pm 0.10) \times 10^{-3}$ .

**$\Gamma(J/\psi(1S) K(1400)^+)/\Gamma(J/\psi(1S) K(1270)^+)$   $\Gamma_{262}/\Gamma_{261}$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.30</b>	90	ABE	01L	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

**$\Gamma(J/\psi(1S) \eta K^+)/\Gamma_{\text{total}}$   $\Gamma_{263}/\Gamma$**

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>12.4±1.4 OUR AVERAGE</b>			
12.7±1.1±1.1	<sup>1</sup> IWASHITA	14	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
10.8±2.3±2.4	<sup>1</sup> AUBERT	04Y	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(X^{c-\text{odd}}(3872) K^+, X^{c-\text{odd}} \rightarrow J/\psi \eta)/\Gamma_{\text{total}}$   $\Gamma_{264}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.8 × 10<sup>-6</sup></b>	90	IWASHITA	14	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

**$\Gamma(\psi(4160) K^+, \psi \rightarrow J/\psi \eta)/\Gamma_{\text{total}}$   $\Gamma_{265}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.4 × 10<sup>-6</sup></b>	90	IWASHITA	14	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

**$\Gamma(J/\psi(1S) \eta' K^+)/\Gamma_{\text{total}}$   $\Gamma_{266}/\Gamma$**

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;8.8</b>	90	<sup>1</sup> XIE	07	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) \phi K^+)/\Gamma_{\text{total}}$   $\Gamma_{267}/\Gamma$**

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>5.0 ± 0.4 OUR AVERAGE</b>			
5.00±0.37±0.15	LEES	15	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
4.4 ± 1.4 ± 0.5	<sup>1</sup> AUBERT	030	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
8.8 $^{+3.5}_{-3.0} \pm 1.3$	<sup>2</sup> ANASTASSOV	00	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ANASTASSOV 00 finds 10 events on a background of  $0.5 \pm 0.2$ . Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ , a uniform Dalitz plot distribution, isotropic  $J/\psi(1S)$  and  $\phi$  decays, and  $B(B^+ \rightarrow J/\psi(1S) \phi K^+) = B(B^0 \rightarrow J/\psi(1S) \phi K^0)$ .

**$\Gamma(X(4140) K^+, X \rightarrow J/\psi(1S) \phi)/\Gamma(J/\psi(1S) \phi K^+)$   $\Gamma_{268}/\Gamma_{267}$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.19 ± 0.07±0.04</b>		<sup>1</sup> ABZOV	14A	D0 $p\bar{p}$ at 1.96 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.133 90 LEES 15 BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<0.07 90 <sup>2</sup> AAIJ 12AA LHCB  $p\bar{p}$  at 7 TeV

<sup>1</sup> Reported a threshold enhancement in the  $J/\psi\phi$  mass distribution consistent with the  $X(4140)$  state with a statistical significance of 3.1 standard deviations.

<sup>2</sup> Branching fractions are normalized to  $382 \pm 22$  events of  $B^+ \rightarrow J/\psi\phi K^+$ .

**$\Gamma(X(4274) K^+, X \rightarrow J/\psi(1S) \phi)/\Gamma(J/\psi(1S) \phi K^+)$   $\Gamma_{269}/\Gamma_{267}$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.08</b>	90	<sup>1</sup> AAIJ	12AA	LHCB $p\bar{p}$ at 7 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.181 90 LEES 15 BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Branching fractions are normalized to  $382 \pm 22$  events of  $B^+ \rightarrow J/\psi\phi K^+$ .

**$\Gamma(J/\psi(1S) \omega K^+)/\Gamma_{\text{total}}$   $\Gamma_{270}/\Gamma$**

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.2±0.1<math>^{+0.6}_{-0.3}</math></b>	<sup>1</sup> DEL-AMO-SA...10B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.5±0.2±0.4 <sup>1</sup> AUBERT 08w BABR Repl. by DEL-AMO-SANCHEZ 10B

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(X(3872) K^+, X \rightarrow J/\psi \omega)/\Gamma_{\text{total}}$   $\Gamma_{271}/\Gamma$**

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>6±2±1</b>	<sup>1</sup> DEL-AMO-SA...10B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(X(3872) K^+, X \rightarrow p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{229}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.7 × 10<sup>-8</sup></b>	95	<sup>1</sup> AAIJ	13S	LHCB $p\bar{p}$ at 7 TeV

<sup>1</sup> Measured relative to  $B^+ \rightarrow J/\psi K^+$  decay with charmonia reconstructed in  $p\bar{p}$  final state and using  $B(B^+ \rightarrow J/\psi K^+) = (1.013 \pm 0.034) \times 10^{-3}$  and  $B(J/\psi \rightarrow p\bar{p}) = (2.17 \pm 0.07) \times 10^{-3}$ .

**$\Gamma(X(3915) K^+, X \rightarrow J/\psi \omega)/\Gamma_{\text{total}}$   $\Gamma_{272}/\Gamma$**

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.0<math>^{+0.7+0.5}_{-0.6-0.3}</math></b>	<sup>1</sup> DEL-AMO-SA...10B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.9 $^{+1.0}_{-0.9} \pm 0.5$  <sup>1</sup> AUBERT 08w BABR Repl. by DEL-AMO-SANCHEZ 10B

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(X(3915) K^+, X \rightarrow p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{259}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.1 × 10<sup>-8</sup></b>	95	<sup>1</sup> AAIJ	13S	LHCB $p\bar{p}$ at 7 TeV

<sup>1</sup> Measured relative to  $B^+ \rightarrow J/\psi K^+$  decay with charmonia reconstructed in  $p\bar{p}$  final state and using  $B(B^+ \rightarrow J/\psi K^+) = (1.013 \pm 0.034) \times 10^{-3}$  and  $B(J/\psi \rightarrow p\bar{p}) = (2.17 \pm 0.07) \times 10^{-3}$ .

**$\Gamma(J/\psi(1S) \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{273}/\Gamma$**

VALUE	DOCUMENT ID	TECN	COMMENT
<b>(4.1±0.4) × 10<sup>-5</sup> OUR FIT</b>			Error includes scale factor of 2.6.
<b>(3.8±0.6±0.3) × 10<sup>-5</sup></b>	<sup>1</sup> ABE	03B	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) \pi^+)/\Gamma(J/\psi(1S) K^+)$   $\Gamma_{273}/\Gamma_{255}$**

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.0 ± 0.4 OUR FIT</b>				Error includes scale factor of 3.3.
<b>4.0 ± 0.4 OUR AVERAGE</b>				Error includes scale factor of 3.2.
3.83±0.11±0.07		AAIJ	12AC	LHCB $p\bar{p}$ at 7 TeV
4.86±0.82±0.15		ABULENCIA	09	CDF $p\bar{p}$ at 1.96 TeV
5.37±0.45±0.11		AUBERT	04P	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
5.0 $^{+1.9}_{-1.7} \pm 0.1$		ABE	96R	CDF $p\bar{p}$ 1.8 TeV
5.2 ± 2.4		BISHAI	96	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.91±0.78±0.19 AUBERT 02F BABR Repl. by AUBERT 04P

4.3 ± 2.3 <sup>5</sup> <sup>1</sup> ALEXANDER 95 CLE2 Sup. by BISHAI 96

<sup>1</sup> Assumes equal production of  $B^+ B^-$  and  $B^0 \bar{B}^0$  on  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) \rho^+)/\Gamma_{\text{total}}$   $\Gamma_{274}/\Gamma$**

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>5.0±0.7±0.3</b>		<sup>1</sup> AUBERT	07AC	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<77 90 BISHAI 96 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) \pi^+ \pi^0 \text{ nonresonant})/\Gamma_{\text{total}}$   $\Gamma_{275}/\Gamma$**

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.73</b>	90	<sup>1</sup> AUBERT	07AC	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

**$\Gamma(J/\psi(1S) a_1(1260)^+)/\Gamma_{\text{total}}$   $\Gamma_{276}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.2 × 10<sup>-3</sup></b>	90	BISHAI	96	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

**$\Gamma(J/\psi p\bar{p}\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{277}/\Gamma$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.0 × 10<sup>-7</sup></b>	90	<sup>1</sup> AAIJ	13Z	LHCB $p\bar{p}$ at 7 TeV

<sup>1</sup> Uses  $B(B_S^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (1.98 \pm 0.20) \times 10^{-4}$ .

# Meson Particle Listings

$B^\pm$

## $\Gamma(J/\psi(1S) \rho \bar{\lambda})/\Gamma_{\text{total}}$ $\Gamma_{278}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>11.8±3.1 OUR AVERAGE</b>				
11.7±2.8 <sup>+1.8</sup> <sub>-2.3</sub>		<sup>1</sup> XIE	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$
12 <sup>+9</sup> <sub>-6</sub>		<sup>1</sup> AUBERT	03k	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<41	90	ZANG	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

## $\Gamma(J/\psi(1S) \Xi^0 \rho)/\Gamma_{\text{total}}$ $\Gamma_{279}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 × 10<sup>-5</sup></b>	90	<sup>1</sup> XIE	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

## $\Gamma(J/\psi(1S) D^+)/\Gamma_{\text{total}}$ $\Gamma_{280}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;12</b>	90	<sup>1</sup> AUBERT	05U	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

## $\Gamma(J/\psi(1S) \bar{D}^0 \pi^+)/\Gamma_{\text{total}}$ $\Gamma_{281}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5</b>	90	<sup>1</sup> ZHANG	05B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.2	90	<sup>1</sup> AUBERT	05R	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

## $\Gamma(\psi(2S) \pi^+)/\Gamma_{\text{total}}$ $\Gamma_{282}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.44±0.22±0.20</b>	<sup>1</sup> BHARDWAJ	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

## $\Gamma(\psi(2S) \pi^+)/\Gamma(\psi(2S) K^+)$ $\Gamma_{282}/\Gamma_{283}$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.97±0.29 OUR AVERAGE</b>			
3.95±0.40±0.12	AAIJ	12Ac	LHCB $pp$ at 7 TeV
3.99±0.36±0.17	BHARDWAJ	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$

## $\Gamma(\psi(2S) K^+)/\Gamma_{\text{total}}$ $\Gamma_{283}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.26± 0.24 OUR FIT</b>				
<b>6.5 ± 0.4 OUR AVERAGE</b>				
6.65± 0.17±0.55		<sup>1</sup> GULER	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
4.9 ± 1.6 ± 0.4		<sup>2</sup> AUBERT	06e	BABR $e^+e^- \rightarrow \Upsilon(4S)$
6.17± 0.32±0.44		<sup>1</sup> AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
7.8 ± 0.7 ± 0.9		<sup>1</sup> RICHICHI	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
18 ± 8 ± 4	5	<sup>1</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.9 ± 0.6		<sup>1</sup> ABE	03B	BELL Repl. by GULER 11
6.4 ± 0.5 ± 0.8		<sup>1</sup> AUBERT	02	BABR Repl. by AUBERT 05J
6.1 ± 2.3 ± 0.9	7	<sup>1</sup> ALAM	94	CLE2 Repl. by RICHICHI 01
<5 at 90% CL		<sup>1</sup> BORTOLETTO	92	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
22 ± 17	3	<sup>3</sup> ALBRECHT	87D	ARG $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Perform measurements of absolute branching fractions using a missing mass technique.  
<sup>3</sup> ALBRECHT 87D assume  $B^+B^-/B^0\bar{B}^0$  ratio is 55/45. Superseded by ALBRECHT 90J.

## $\Gamma(\psi(2S) K^+)/\Gamma(J/\psi(1S) K^+)$ $\Gamma_{283}/\Gamma_{255}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.610±0.019 OUR FIT</b>			
<b>0.603±0.021 OUR AVERAGE</b>			
0.59 ± 0.11 ± 0.02	<sup>1</sup> AAIJ	13s	LHCB $pp$ at 7 TeV
0.604±0.018±0.013	<sup>2,3</sup> AAIJ	12L	LHCB $pp$ at 7 TeV
0.63 ± 0.05 ± 0.08	ABAZOV	09Y	D0 $p\bar{p}$ at 1.96 TeV
0.558±0.082±0.056	ABE	98O	CDF $p\bar{p}$ 1.8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.64 ± 0.06 ± 0.07	<sup>4</sup> AUBERT	02	BABR $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> AAIJ 13s reports  $[\Gamma(B^+ \rightarrow \psi(2S) K^+)/\Gamma(B^+ \rightarrow J/\psi(1S) K^+)] \times [B(\psi(2S) \rightarrow p\bar{p})] / [B(J/\psi(1S) \rightarrow p\bar{p})] = 0.080 \pm 0.012 \pm 0.009$  which we multiply or divide by our best values  $B(\psi(2S) \rightarrow p\bar{p}) = (2.88 \pm 0.09) \times 10^{-4}$ ,  $B(J/\psi(1S) \rightarrow p\bar{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.  
<sup>2</sup> AAIJ 12L reports  $0.594 \pm 0.006 \pm 0.016 \pm 0.015$  from a measurement of  $[\Gamma(B^+ \rightarrow \psi(2S) K^+)/\Gamma(B^+ \rightarrow J/\psi(1S) K^+)] \times [B(J/\psi(1S) \rightarrow e^+e^-)] / [B(\psi(2S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.94 \pm 0.06) \times 10^{-2}$ ,  $B(\psi(2S) \rightarrow e^+e^-) = (7.72 \pm 0.17) \times 10^{-3}$ , which we rescale to our best values  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ ,  $B(\psi(2S) \rightarrow e^+e^-) = (7.89 \pm 0.17) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.  
<sup>3</sup> Assumes  $B(J/\psi \rightarrow \mu^+\mu^-) / B(\psi(2S) \rightarrow \mu^+\mu^-) = B(J/\psi \rightarrow e^+e^-) / B(\psi(2S) \rightarrow e^+e^-) = 7.69 \pm 0.19$ .  
<sup>4</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\psi(2S) K^*(892)^+)/\Gamma_{\text{total}}$ $\Gamma_{284}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>6.7 ± 1.4 OUR AVERAGE</b>				Error includes scale factor of 1.3.
5.92±0.85±0.89		<sup>1</sup> AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
9.2 ± 1.9 ± 1.2		<sup>1</sup> RICHICHI	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<30	90	<sup>1</sup> ALAM	94	CLE2 Repl. by RICHICHI 01
<35	90	<sup>1</sup> BORTOLETTO	092	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
<49	90	<sup>1</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

## $\Gamma(\psi(2S) K^*(892)^+)/\Gamma(\psi(2S) K^+)$ $\Gamma_{284}/\Gamma_{283}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.96±0.15±0.09</b>	AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$

## $\Gamma(\psi(2S) K^0 \pi^+)/\Gamma_{\text{total}}$ $\Gamma_{285}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.588±0.034	<sup>1</sup> AUBERT	09AA	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Does not report systematic uncertainties.			

## $\Gamma(\psi(2S) K^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{286}/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.3 ± 0.5 OUR AVERAGE</b>				
4.31± 0.20±0.50		<sup>1</sup> GULER	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
19 ± 11 ± 4	3	<sup>1</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

## $\Gamma(\psi(3770) K^+)/\Gamma_{\text{total}}$ $\Gamma_{287}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.49±0.13 OUR AVERAGE</b>			
3.5 ± 2.5 ± 0.3	<sup>1</sup> AUBERT	06E	BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.48±0.11±0.07	<sup>2</sup> CHISTOV	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Perform measurements of absolute branching fractions using a missing mass technique. <sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

## $\Gamma(\psi(3770) K^+, \psi \rightarrow D^0 \bar{D}^0)/\Gamma_{\text{total}}$ $\Gamma_{288}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.5 ± 0.5 OUR AVERAGE</b>			Error includes scale factor of 1.4.
1.18±0.41±0.15	<sup>1</sup> LEES	15c	BABR $e^+e^- \rightarrow \Upsilon(4S)$
2.2 ± 0.5 ± 0.3	<sup>1</sup> BRODZICKA	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.41±0.30±0.22	<sup>1</sup> AUBERT	08B	BABR Repl. by LEES 15c
3.4 ± 0.8 ± 0.5	<sup>1</sup> CHISTOV	04	BELL Repl. by BRODZICKA 08
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

## $\Gamma(\psi(3770) K^+, \psi \rightarrow D^+ D^-)/\Gamma_{\text{total}}$ $\Gamma_{289}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.94±0.35 OUR AVERAGE</b>			
0.84±0.32±0.21	<sup>1</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.4 ± 0.8 ± 0.2	<sup>1</sup> CHISTOV	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

## $\Gamma(\psi(4040) K^+)/\Gamma_{\text{total}}$ $\Gamma_{290}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.3 × 10<sup>-4</sup></b>	90	AAIJ	13Bc	LHCB $pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0 × 10 <sup>-3</sup>	90	<sup>1</sup> IWASHITA	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> IWASHITA 14 reports $[\Gamma(B^+ \rightarrow \psi(4040) K^+)/\Gamma_{\text{total}}] \times [B(\psi(4040) \rightarrow J/\psi \eta)] < 15.5 \times 10^{-6}$ which we divide by our best value $B(\psi(4040) \rightarrow J/\psi \eta) = 5.2 \times 10^{-3}$ .				

## $\Gamma(\psi(4160) K^+)/\Gamma_{\text{total}}$ $\Gamma_{291}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>5.1<sup>+1.3+2.5</sup><sub>-1.2-2.4</sub></b>	<sup>1</sup> AAIJ	13Bc	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 13Bc reports $[\Gamma(B^+ \rightarrow \psi(4160) K^+)/\Gamma_{\text{total}}] \times B(\psi(4160) \rightarrow \mu^+\mu^-) = (3.5^{+0.9}_{-0.8}) \times 10^{-9}$ which we devide by our best value $B(\psi(4160) \rightarrow e^+e^-) = (6.9 \pm 3.3) \times 10^{-6}$ assuming lepton universality. Our first error is their experiment's error and our second error is the systematic error from using our best value.			

## $\Gamma(\psi(4160) K^+, \psi \rightarrow \bar{D}^0 D^0)/\Gamma_{\text{total}}$ $\Gamma_{292}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.84±0.41±0.33</b>	<sup>1</sup> LEES	15c	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

## $\Gamma(\chi_{c0} \pi^+, \chi_{c0} \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{293}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.1</b>	90	<sup>1</sup> AUBERT	09L	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.3	90	<sup>1</sup> AUBERT,B	05G	BABR Repl. by AUBERT 09L
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\chi_{c0}(1P)K^+)/\Gamma_{\text{total}}$		$\Gamma_{294}/\Gamma$		
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**1.50<sup>+0.15</sup><sub>-0.14</sub> OUR AVERAGE**

1.84 $\pm$ 0.25 $\pm$ 0.14	1,2	LEES	12o	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.68 $\pm$ 0.32 $\pm$ 0.16	1,3	LEES	12o	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.8 $\pm$ 0.9 $\pm$ 0.1	4	LEES	11i	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.26 $\pm$ 0.28 $\pm$ 0.05 -0.25	1,5	AUBERT	08Ai	BABR $e^+e^- \rightarrow \Upsilon(4S)$
4.8 $\pm$ 2.2 $\pm$ 0.2	6	AUBERT,BE	06M	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.12 $\pm$ 0.12 $\pm$ 0.30 -0.20	1	GARMASH	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.7	95	7 AAIJ	13s	LHCB $pp$ at 7 TeV
<5	90	1,8 WICHT	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<1.8	90	9 AUBERT	06E	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.84 $\pm$ 0.32 $\pm$ 0.31	1,10	AUBERT	06o	BABR Repl. by LEES 12o
<8.9	90	1 AUBERT	05K	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.39 $\pm$ 0.49 $\pm$ 0.11	11	AUBERT,B	05N	BABR Repl. by AUBERT 08Ai
1.96 $\pm$ 0.35 $\pm$ 2.00 -0.42	1	GARMASH	05	BELL Repl. by GARMASH 06
2.7 $\pm$ 0.7	12	AUBERT	04T	BABR Repl. by AUBERT,B 04P
3.0 $\pm$ 0.8 $\pm$ 0.3	13	AUBERT,B	04P	BABR Repl. by AUBERT,B 05N
6.0 $\pm$ 2.1 $\pm$ 1.1 -1.8	14	ABE	02B	BELL Repl. by GARMASH 05
<4.8	90	15 EDWARDS	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Measured in the  $B^+ \rightarrow K^+K^-K^+$  decay.

<sup>3</sup> Measured in the  $B^+ \rightarrow K^+K_S^0 K_S^0$  decay.

<sup>4</sup> LEES 11i reports  $[\Gamma(B^+ \rightarrow \chi_{c0}(1P)K^+)/\Gamma_{\text{total}}] \times [B(\chi_{c0}(1P) \rightarrow \pi\pi)] = (1.53 \pm 0.66 \pm 0.27) \times 10^{-6}$  which we divide by our best value  $B(\chi_{c0}(1P) \rightarrow \pi\pi) = (8.33 \pm 0.35) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>5</sup> AUBERT 08Ai reports  $(0.70 \pm 0.10 \pm 0.12 \pm 0.10) \times 10^{-6}$  for  $B(B^+ \rightarrow \chi_{c0}K^+) \times B(\chi_{c0} \rightarrow \pi^+\pi^-)$ . We compute  $B(B^+ \rightarrow \chi_{c0}K^+)$  using the PDG value  $B(\chi_{c0} \rightarrow \pi\pi) = (8.33 \pm 0.35) \times 10^{-3}$  and 2/3 for the  $\pi^+\pi^-$  fraction. Our first error is their experiment's error and the second error is systematic error from using our best value.

<sup>6</sup> AUBERT,BE 06M reports  $[\Gamma(B^+ \rightarrow \chi_{c0}(1P)K^+)/\Gamma_{\text{total}}] \times [B(\chi_{c0}(1P) \rightarrow \gamma J/\psi(1S))] = (6.1 \pm 2.6 \pm 1.1) \times 10^{-6}$  which we divide by our best value  $B(\chi_{c0}(1P) \rightarrow \gamma J/\psi(1S)) = (1.27 \pm 0.06) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. The significance of the observed signal is 2.4  $\sigma$ .

<sup>7</sup> AAIJ 13s reports  $[\Gamma(B^+ \rightarrow \chi_{c0}(1P)K^+)/\Gamma_{\text{total}}] \times [B(\chi_{c0}(1P) \rightarrow p\bar{p})] < 6 \times 10^{-8}$  which we divide by our best value  $B(\chi_{c0}(1P) \rightarrow p\bar{p}) = 2.25 \times 10^{-4}$ .

<sup>8</sup> WICHT 08 reports  $[\Gamma(B^+ \rightarrow \chi_{c0}(1P)K^+)/\Gamma_{\text{total}}] \times [B(\chi_{c0}(1P) \rightarrow \gamma\gamma)] < 0.11 \times 10^{-6}$  which we divide by our best value  $B(\chi_{c0}(1P) \rightarrow \gamma\gamma) = 2.23 \times 10^{-4}$ .

<sup>9</sup> Perform measurements of absolute branching fractions using a missing mass technique.

<sup>10</sup> Measured in the  $B^+ \rightarrow K^+K^-K^+$  decay.

<sup>11</sup> AUBERT,B 05N reports  $(0.66 \pm 0.22 \pm 0.08) \times 10^{-6}$  for  $B(B^+ \rightarrow \chi_{c0}^0 K^+) \times B(\chi_{c0}^0 \rightarrow \pi^+\pi^-)$ . We compute  $B(B^+ \rightarrow \chi_{c0}^0 K^+)$  using the PDG value  $B(\chi_{c0}^0 \rightarrow \pi^+\pi^-) = (7.1 \pm 0.6) \times 10^{-3}$  and 2/3 for the  $\pi^+\pi^-$  fraction.

<sup>12</sup> The measurement performed using decay channels  $\chi_{c0} \rightarrow \pi^+\pi^-$  and  $\chi_{c0} \rightarrow K^+K^-$ . The ratio of the branching ratios for these channels is found to be consistent with world average.

<sup>13</sup> AUBERT 04P reports  $B(B^+ \rightarrow \chi_{c0}K^+) \times B(\chi_{c0} \rightarrow \pi^+\pi^-) = (1.5 \pm 0.4 \pm 0.1) \times 10^{-6}$  and used PDG value of  $B(\chi_{c0} \rightarrow \pi\pi) = (7.4 \pm 0.8) \times 10^{-3}$  and Clebsh-Gordan coefficient to compute  $B(B^\pm \rightarrow \chi_{c0}K^+)$ .

<sup>14</sup> ABE 02B measures the ratio of  $B(B^+ \rightarrow \chi_{c0}K^+)/B(B^+ \rightarrow J/\psi(1S)K^+) = 0.60 \pm 0.21 - 0.18 \pm 0.05 \pm 0.08$ , where the third error is due to the uncertainty in the  $B(\chi_{c0} \rightarrow \pi^+\pi^-)$ , and uses  $B(B^+ \rightarrow J/\psi(1S)K^+) = (10.0 \pm 1.0) \times 10^{-4}$  to obtain the result.

<sup>15</sup> EDWARDS 01 assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ . The correlated uncertainties (28.3)% from  $B(J/\psi(1S) \rightarrow \eta\eta_c)$  in those modes have been accounted for.

$\Gamma(\chi_{c0}K^*(892)^+)/\Gamma_{\text{total}}$		$\Gamma_{295}/\Gamma$		
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**< 2.1**

<2.6	90	1 AUBERT	08BD	BABR Repl. by AUBERT 08BD
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c2}\pi^+, \chi_{c2} \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{296}/\Gamma$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**<0.1**

<0.1	90	1 AUBERT	09L	BABR $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c2}K^+)/\Gamma_{\text{total}}$		$\Gamma_{297}/\Gamma$		
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**1.11<sup>+0.36</sup><sub>-0.34</sub>  $\pm$  0.09**

< 1.8	90	2 AUBERT	09B	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<20	90	3 AUBERT	06E	BABR $e^+e^- \rightarrow \Upsilon(4S)$
< 2.9	90	1 SONI	06	BELL Repl. by BHARDWAJ 11
< 3.0	90	1 AUBERT	05K	BABR Repl. by AUBERT 06E

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses  $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .

<sup>3</sup> Perform measurements of absolute branching fractions using a missing mass technique.

$\Gamma(B^+ \rightarrow \chi_{c2}K^+)/\Gamma_{\text{total}} \times \Gamma(\chi_{c2}(1P) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}$		$\Gamma_{297}/\Gamma \times \Gamma_{\chi_{c2}(1P)}^{\chi_{c2}(1P)}/\Gamma_{\chi_{c2}(1P)}$		
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**<0.09**

<0.09	90	1 WICHT	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c2}K^*(892)^+)/\Gamma_{\text{total}}$		$\Gamma_{298}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**<12  $\times 10^{-5}$**

<12.7 $\times 10^{-5}$	90	2 SONI	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
< 1.2 $\times 10^{-5}$	90	2 AUBERT	05K	BABR Repl. by AUBERT 09B

<sup>1</sup> Uses  $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1}(1P)\pi^+)/\Gamma_{\text{total}}$		$\Gamma_{299}/\Gamma$		
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**2.2 $\pm$ 0.4 $\pm$ 0.3**

2.2 $\pm$ 0.4 $\pm$ 0.3	1	KUMAR	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1}(1P)K^+)/\Gamma_{\text{total}}$		$\Gamma_{300}/\Gamma$		
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT

**4.79 $\pm$  0.23 OUR AVERAGE**

4.94 $\pm$ 0.11 $\pm$ 0.33	1	BHARDWAJ	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
4.5 $\pm$ 0.1 $\pm$ 0.3	2	AUBERT	09B	BABR $e^+e^- \rightarrow \Upsilon(4S)$
8.1 $\pm$ 1.4 $\pm$ 0.7	3	AUBERT	06E	BABR $e^+e^- \rightarrow \Upsilon(4S)$
15.5 $\pm$ 5.4 $\pm$ 2.0	4	ACOSTA	02F	CDF $p\bar{p}$ 1.8 TeV
5.2 $\pm$ 0.4 $\pm$ 0.2	5	AUBERT,BE	06M	BABR Repl. by AUBERT 09B
4.49 $\pm$ 0.19 $\pm$ 0.53	1	SONI	06	BELL Repl. by BHARDWAJ 11
5.79 $\pm$ 0.26 $\pm$ 0.65	1	AUBERT	05J	BABR Repl. by AUBERT,BE 06M
6.0 $\pm$ 0.9 $\pm$ 0.2	6	AUBERT	02	BABR Repl. by AUBERT 05J
9.7 $\pm$ 4.0 $\pm$ 0.9	6	1 ALAM	94	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
19 $\pm$ 13 $\pm$ 6	7	ALBRECHT	92E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses  $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .

<sup>3</sup> Perform measurements of absolute branching fractions using a missing mass technique.

<sup>4</sup> ACOSTA 02F uses as reference of  $B(B \rightarrow J/\psi(1S)K^+) \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))] = (10.1 \pm 0.6) \times 10^{-4}$ . The second error includes the systematic error and the uncertainties of the branching ratio.

<sup>5</sup> AUBERT,BE 06M reports  $[\Gamma(B^+ \rightarrow \chi_{c1}(1P)K^+)/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))] = (1.76 \pm 0.07 \pm 0.12) \times 10^{-4}$  which we divide by our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>6</sup> AUBERT 02 reports  $(7.5 \pm 0.9 \pm 0.8) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \chi_{c1}(1P)K^+)/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>7</sup> ALBRECHT 92E assumes no  $\chi_{c2}(1P)$  production and  $B(\Upsilon(4S) \rightarrow B^+B^-) = 50\%$ .

$\Gamma(\chi_{c1}(1P)K^+)/\Gamma(J/\psi(1S)K^+)$		$\Gamma_{300}/\Gamma_{255}$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**0.60 $\pm$ 0.07 $\pm$ 0.02**

0.60 $\pm$ 0.07 $\pm$ 0.02	1	AUBERT	02	BABR $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> AUBERT 02 reports  $0.75 \pm 0.08 \pm 0.05$  from a measurement of  $[\Gamma(B^+ \rightarrow \chi_{c1}(1P)K^+)/\Gamma(B^+ \rightarrow J/\psi(1S)K^+)] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1}(1P)\pi^+)/\Gamma(\chi_{c1}(1P)K^+)$		$\Gamma_{299}/\Gamma_{300}$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**0.043 $\pm$ 0.008 $\pm$ 0.003**

0.043 $\pm$ 0.008 $\pm$ 0.003	1	KUMAR	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1}(1P)K^0\pi^+)/\Gamma(J/\psi(1S)K^0\pi^+)$		$\Gamma_{301}/\Gamma_{256}$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**0.508 $\pm$ 0.030 $\pm$ 0.018**

0.508 $\pm$ 0.030 $\pm$ 0.018	1	LEES	12B	BABR $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> LEES 12B reports  $0.501 \pm 0.024 \pm 0.028$  from a measurement of  $[\Gamma(B^+ \rightarrow \chi_{c1}(1P)K^0\pi^+)/\Gamma(B^+ \rightarrow J/\psi(1S)K^0\pi^+)] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (34.4 \pm 1.5) \times 10^{-2}$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

$B^\pm$

$\Gamma(\chi_{c1}(1P)K^*(892)^+)/\Gamma_{\text{total}}$					$\Gamma_{302}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>3.0 ± 0.6 OUR AVERAGE</b>		Error includes scale factor of 1.1.			
2.6 ± 0.5 ± 0.4		<sup>1</sup> AUBERT	09b	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
4.05 ± 0.59 ± 0.95		<sup>2</sup> SONI	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.94 ± 0.95 ± 0.98		<sup>2</sup> AUBERT	05j	BABR Repl. by AUBERT 09b	
<21	90	<sup>2</sup> ALAM	94	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .					
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\chi_{c1}(1P)K^*(892)^+)/\Gamma(\chi_{c1}(1P)K^+)$					$\Gamma_{302}/\Gamma_{300}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.51 ± 0.17 ± 0.16</b>		AUBERT	05j	BABR $e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(h_c(1P)K^+)/\Gamma_{\text{total}}$					$\Gamma_{303}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.8</b>	90	<sup>1</sup> FANG	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ and $B(h_c \rightarrow \eta_c\gamma) = 50\%$ .					

$\Gamma(h_c(1P)K^+, h_c \rightarrow p\bar{p})/\Gamma_{\text{total}}$					$\Gamma_{304}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;6.4 × 10<sup>-8</sup></b>		<sup>1</sup> AAIJ	13s	LHCB $p\bar{p}$ at 7 TeV	
<sup>1</sup> Measured relative to $B^+ \rightarrow J/\psi K^+$ decay with charmonia reconstructed in $p\bar{p}$ final state and using $B(B^+ \rightarrow J/\psi K^+) = (1.013 \pm 0.034) \times 10^{-3}$ and $B(J/\psi \rightarrow p\bar{p}) = (2.17 \pm 0.07) \times 10^{-3}$ .					

$\Gamma(K^0\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{305}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>23.7 ± 0.8 OUR FIT</b>					
<b>23.8 ± 0.7 OUR AVERAGE</b>					
23.97 ± 0.53 ± 0.71		<sup>1</sup> DUH	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
23.9 ± 1.1 ± 1.0		<sup>1</sup> AUBERT,BE	06c	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
18.8 ± 3.7 ± 2.1		<sup>1</sup> BORNHEIM	03	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
18.8 ± 3.3 ± 1.8					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
22.8 ± 0.8 ± 1.3		<sup>1</sup> LIN	07	BELL Repl. by DUH 13	
26.0 ± 1.3 ± 1.0		<sup>1</sup> AUBERT,BE	05E	BABR Repl. by AUBERT,BE 06c	
22.3 ± 1.7 ± 1.1		<sup>1</sup> AUBERT	04M	BABR Repl. by AUBERT,BE 05E	
22.0 ± 1.9 ± 1.1		<sup>1</sup> CHAO	04	BELL Repl. by LIN 07	
19.4 ± 3.1 ± 3.0		<sup>1</sup> CASEY	02	BELL Repl. by CHAO 04	
13.7 ± 5.7 ± 1.9		<sup>1</sup> ABE	01H	BELL Repl. by CASEY 02	
18.2 ± 3.3 ± 3.0		<sup>1</sup> AUBERT	01E	BABR Repl. by AUBERT 04M	
18.2 ± 4.6 ± 1.6		<sup>1</sup> CRONIN-HEN..00	CLE2	Repl. by BORNHEIM 03	
23 ± 11 ± 10		GODANG	98	CLE2 Repl. by CRONIN-HENNESSY 00	
< 48	90	ASNER	96	CLE2 Repl. by GODANG 98	
<190	90	ALBRECHT	91B	ARG $e^+e^- \rightarrow \Upsilon(4S)$	
<100	90	2 AVERY	89B	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	
<680	90	AVERY	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AVERY 89b reports  $< 9 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{306}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>12.9 ± 0.5 OUR AVERAGE</b>					
12.62 ± 0.31 ± 0.56		<sup>1</sup> DUH	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
13.6 ± 0.6 ± 0.7		<sup>1</sup> AUBERT	07Bc	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
12.9 ± 2.4 ± 1.2		<sup>1</sup> BORNHEIM	03	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
12.9 ± 2.2 ± 1.1					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
12.4 ± 0.5 ± 0.6		<sup>1</sup> LIN	07A	BELL Repl. by DUH 13	
12.0 ± 0.7 ± 0.6		<sup>1</sup> AUBERT	05L	BABR Repl. by AUBERT 07Bc	
12.0 ± 1.3 ± 1.3		<sup>1</sup> CHAO	04	BELL Repl. by LIN 07A	
12.8 ± 1.2 ± 1.1		<sup>1</sup> AUBERT	03L	BABR Repl. by AUBERT 05L	
13.0 ± 2.5 ± 1.3		<sup>1</sup> CASEY	02	BELL Repl. by CHAO 04	
16.3 ± 3.5 ± 1.6		<sup>1</sup> ABE	01H	BELL Repl. by CASEY 02	
10.8 ± 2.1 ± 1.9		<sup>1</sup> AUBERT	01E	BABR Repl. by AUBERT 03L	
11.6 ± 3.0 ± 1.4		<sup>1</sup> CRONIN-HEN..00	CLE2	Repl. by BORNHEIM 03	
<16	90	GODANG	98	CLE2 Repl. by CRONIN-HENNESSY 00	
<14	90	ASNER	96	CLE2 Repl. by GODANG 98	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+\pi^0)/\Gamma(K^0\pi^+)$				$\Gamma_{306}/\Gamma_{305}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.54 ± 0.03 ± 0.04</b>	LIN	07A	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.38 ± 0.98 ± 0.39	ABE	01H	BELL Repl. by LIN 07A	
-1.10 - 0.26				

$\Gamma(\eta'K^+)/\Gamma_{\text{total}}$				$\Gamma_{307}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b>70.6 ± 2.5 OUR AVERAGE</b>				
71.5 ± 1.3 ± 3.2	<sup>1</sup> AUBERT	09AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
63 ± 10 ± 9	<sup>1,2</sup> WICHT	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
69.2 ± 2.2 ± 3.7	<sup>1</sup> SCHUEMANN	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
80 ± 10 ± 9	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
70.0 ± 1.5 ± 2.8	<sup>1</sup> AUBERT	07AE	BABR Repl. by AUBERT 09AV	
68.9 ± 2.0 ± 3.2	<sup>1</sup> AUBERT	05M	BABR Repl. by AUBERT 07AE	
76.9 ± 3.5 ± 4.4	<sup>1</sup> AUBERT	03W	BABR Repl. by AUBERT 05M	
79 ± 12 ± 11	<sup>1</sup> ABE	01M	BELL Repl. by SCHUEMANN 06	
70 ± 8 ± 5	<sup>1</sup> AUBERT	01G	BABR Repl. by AUBERT 03W	
65 ± 15 ± 14	BEHRENS	98	CLE2 Repl. by RICHICHI 00	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> WICHT 08 reports  $[\Gamma(B^+ \rightarrow \eta'K^+)/\Gamma_{\text{total}}] \times [B(\eta'(958) \rightarrow \gamma\gamma)] = (1.40 \pm 0.16 \pm 0.15 \pm 0.15) \times 10^{-6}$  which we divide by our best value  $B(\eta'(958) \rightarrow \gamma\gamma) = (2.21 \pm 0.08) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\eta'K^*(892)^+)/\Gamma_{\text{total}}$					$\Gamma_{308}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>4.8 ± 1.6 ± 1.4 ± 0.8</b>		<sup>1</sup> DEL-AMO-SA..10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.9 ± 1.9 ± 1.7 ± 0.8		<sup>1</sup> AUBERT	07E	BABR Repl. by DEL-AMO-SANCHEZ 10A	
< 2.9	90	<sup>1</sup> SCHUEMANN	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<14	90	<sup>1</sup> AUBERT,B	04D	BABR Repl. by AUBERT 07E	
<35	90	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
<13	90	BEHRENS	98	CLE2 Repl. by RICHICHI 00	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\eta'K_s^*(1430)^+)/\Gamma_{\text{total}}$					$\Gamma_{309}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b>5.2 ± 1.9 ± 1.0</b>	<sup>1</sup> DEL-AMO-SA..10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$		
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\eta'K_s^*(1430)^+)/\Gamma_{\text{total}}$					$\Gamma_{310}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b>28.0 ± 4.6 ± 4.3 ± 2.6</b>	<sup>1</sup> DEL-AMO-SA..10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$		
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\eta K^+)/\Gamma_{\text{total}}$					$\Gamma_{311}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>2.4 ± 0.4 OUR AVERAGE</b>		Error includes scale factor of 1.7.			
2.12 ± 0.23 ± 0.11		<sup>1</sup> HOI	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
2.94 ± 0.39 ± 0.34 ± 0.21		<sup>1</sup> AUBERT	09AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
2.2 ± 2.8 ± 2.2		<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.21 ± 0.48 ± 0.42 ± 0.01		<sup>1,2</sup> WICHT	08	BELL Repl. by HOI 12	
3.7 ± 0.4 ± 0.1		<sup>1</sup> AUBERT	07AE	BABR Repl. by AUBERT 09AV	
1.9 ± 0.3 ± 0.2 ± 0.1		<sup>1</sup> CHANG	07B	BELL Repl. by HOI 12	
3.3 ± 0.6 ± 0.3		<sup>1</sup> AUBERT,B	05K	BABR Repl. by AUBERT 07AE	
2.1 ± 0.6 ± 0.2		<sup>1</sup> CHANG	05A	BELL Repl. by CHANG 07B	
3.4 ± 0.8 ± 0.2		<sup>1</sup> AUBERT	04H	BABR Repl. by AUBERT,B 05K	
<14	90	BEHRENS	98	CLE2 Repl. by RICHICHI 00	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>2</sup> WICHT 08 reports $[\Gamma(B^+ \rightarrow \eta K^+)/\Gamma_{\text{total}}] \times [B(\eta \rightarrow 2\gamma)] = (0.87 \pm 0.16 \pm 0.10 \pm 0.15 \pm 0.07) \times 10^{-6}$ which we divide by our best value $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\eta K^*(892)^+)/\Gamma_{\text{total}}$					$\Gamma_{312}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>19.3 ± 1.6 OUR AVERAGE</b>					
19.3 ± 2.0 ± 1.5		<sup>1</sup> WANG	07B	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
18.9 ± 1.8 ± 1.3		<sup>1</sup> AUBERT,B	06H	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
26.4 ± 9.6 ± 8.2 ± 3.3		<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					



See key on page 601

## Meson Particle Listings

 $B^\pm$ 

$25.6 \pm 4.0 \pm 2.4$	<sup>1</sup> AUBERT,B	04D	BABR	Repl. by AUBERT,B 06H Repl. by RICHICHI 00
<30	90	BEHRENS	98	CLE2
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\eta K_0^*(1430)^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{313}/\Gamma$
VALUE (units $10^{-6}$ )				
$18.2 \pm 2.6 \pm 2.6$	<sup>1</sup> AUBERT,B	06H	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\eta K_2^*(1430)^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{314}/\Gamma$
VALUE (units $10^{-6}$ )				
$9.1 \pm 2.7 \pm 1.4$	<sup>1</sup> AUBERT,B	06H	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\eta(1295) K^+ \times B(\eta(1295) \rightarrow \eta\pi\pi))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{315}/\Gamma$
VALUE (units $10^{-6}$ )				
$2.9 \pm 0.8 \pm 0.2$	<sup>1</sup> AUBERT	08x	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\eta(1405) K^+ \times B(\eta(1405) \rightarrow \eta\pi\pi))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{316}/\Gamma$
VALUE (units $10^{-6}$ )				
<1.3	90	<sup>1</sup> AUBERT	08x BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\eta(1405) K^+ \times B(\eta(1405) \rightarrow K^* K))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{317}/\Gamma$
VALUE (units $10^{-6}$ )				
<1.2	90	<sup>1</sup> AUBERT	08x BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\eta(1475) K^+ \times B(\eta(1475) \rightarrow K^* K))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{318}/\Gamma$
VALUE (units $10^{-6}$ )				
$13.8 \pm 1.8 \pm 1.0$ $-1.7 - 0.6$	<sup>1</sup> AUBERT	08x	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\tilde{\eta}_1(1285) K^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{319}/\Gamma$
VALUE (units $10^{-6}$ )				
<2.0	90	<sup>1</sup> AUBERT	08x BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\tilde{\eta}_1(1420) K^+ \times B(\tilde{\eta}_1(1420) \rightarrow \eta\pi\pi))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{320}/\Gamma$
VALUE (units $10^{-6}$ )				
<2.9	90	<sup>1</sup> AUBERT	08x BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\tilde{\eta}_1(1420) K^+ \times B(\tilde{\eta}_1(1420) \rightarrow K^* K))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{321}/\Gamma$
VALUE (units $10^{-6}$ )				
<4.1	90	<sup>1</sup> AUBERT	08x BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\phi(1680) K^+ \times B(\phi(1680) \rightarrow K^* K))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{322}/\Gamma$
VALUE (units $10^{-6}$ )				
<3.4	90	<sup>1</sup> AUBERT	08x BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(f_0(1500) K^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{323}/\Gamma$
VALUE (units $10^{-6}$ )				
<b>3.7 <math>\pm</math> 2.2 OUR AVERAGE</b>				
$17 \pm 4 \pm 12$	<sup>1</sup> LEES	12o	BABR $e^+e^- \rightarrow T(4S)$	
$20 \pm 10 \pm 27$	<sup>2</sup> LEES	12o	BABR $e^+e^- \rightarrow T(4S)$	
$3.1 \pm 2.3 \pm 0.2$	<sup>3,4</sup> AUBERT	08A1	BABR $e^+e^- \rightarrow T(4S)$	

- • • We do not use the following data for averages, fits, limits, etc. • • •
- <19 90 <sup>4,5</sup> AUBERT,B 05N BABR Repl. by AUBERT 08A1
- <sup>1</sup> Measured in the  $B^+ \rightarrow K^+ K^- K^+$  decay.
- <sup>2</sup> Measured in the  $B^+ \rightarrow K^+ K_S^0 K^0$  decay.
- <sup>3</sup> AUBERT 08A1 reports  $B(B^+ \rightarrow f_0(1500) K^+) \cdot B(f_0(1500) \rightarrow \pi^+ \pi^-) = (0.73 \pm 0.21 \pm 0.47) \times 10^{-6}$ . We divide this result by our best value of  $B(f_0(1500) \rightarrow \pi\pi) = (34.9 \pm 2.3) \times 10^{-2}$  multiplied by 2/3 to account for the  $\pi^+ \pi^-$  fraction. Our first quoted uncertainty is the combined experiment's uncertainty and our second is the systematic uncertainty from using out best value.
- <sup>4</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $T(4S)$ .
- <sup>5</sup> AUBERT,B 05N reports  $B(B^+ \rightarrow f_0(1500) K^+) \cdot B(f_0(1500) \rightarrow \pi^+ \pi^-) < 4.4 \times 10^{-6}$ . We divide this result by our best value of  $B(f_0(1500) \rightarrow \pi\pi) = (34.9 \pm 2.3) \times 10^{-2}$  multiplied by 2/3 to account for the  $\pi^+ \pi^-$  fraction. Our first quoted uncertainty is the combined experiment's uncertainty and our second is the systematic uncertainty from using out best value.

$\Gamma(\omega K^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{324}/\Gamma$
VALUE (units $10^{-6}$ )				
<b>6.5 <math>\pm</math> 0.4 OUR AVERAGE</b>				
$6.8 \pm 0.4 \pm 0.4$	<sup>1</sup> CHOBANOVA	14	BELL $e^+e^- \rightarrow T(4S)$	
$6.3 \pm 0.5 \pm 0.3$	<sup>1</sup> AUBERT	07AE	BABR $e^+e^- \rightarrow T(4S)$	
$3.2 \pm 2.4 \pm 1.9$	<sup>1</sup> JESSOP	00	CLE2 $e^+e^- \rightarrow T(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$6.1 \pm 0.6 \pm 0.4$	<sup>1</sup> AUBERT,B	06E	BABR AUBERT 07AE	
$8.1 \pm 0.6 \pm 0.6$	<sup>1</sup> JEN	06	BELL Repl. by CHOBANOVA 14	
$4.8 \pm 0.8 \pm 0.4$	<sup>1</sup> AUBERT	04H	BABR Repl. by AUBERT,B 06E	
$6.5 \pm 1.3 \pm 1.2$	<sup>1</sup> WANG	04A	BELL Repl. by JEN 06	
$9.2 \pm 2.6 \pm 2.3$	<sup>1</sup> LU	02	BELL Repl. by WANG 04A	
<4	90	<sup>1</sup> AUBERT	01G BABR $e^+e^- \rightarrow T(4S)$	
$1.5 \pm 0.7 \pm 0.2$	<sup>1</sup> BERGFELD	98	CLE2 Repl. by JESSOP 00	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\omega K^*(892)^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{325}/\Gamma$
VALUE (units $10^{-6}$ )				
< 7.4	90	<sup>1</sup> AUBERT	09H BABR $e^+e^- \rightarrow T(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 3.4	90	<sup>1</sup> AUBERT,B	06T BABR Repl. by AUBERT 09H	
< 7.4	90	<sup>1</sup> AUBERT	05o BABR Repl. by AUBERT,B 06T	
<87	90	<sup>1</sup> BERGFELD	98 CLE2	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\omega(K\pi)_0^{*+})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{326}/\Gamma$
VALUE (units $10^{-6}$ )				
$27.5 \pm 3.0 \pm 2.6$	<sup>1</sup> AUBERT	09H	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\omega K_0^*(1430)^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{327}/\Gamma$
VALUE (units $10^{-6}$ )				
<b>24.0 <math>\pm</math> 2.6 <math>\pm</math> 4.4</b>	<sup>1</sup> AUBERT	09H	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(\omega K_2^*(1430)^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{328}/\Gamma$
VALUE (units $10^{-6}$ )				
<b>21.5 <math>\pm</math> 3.6 <math>\pm</math> 2.4</b>	<sup>1</sup> AUBERT	09H	BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $T(4S)$ .				

$\Gamma(a_0(980)^0 K^+ \times B(a_0(980)^0 \rightarrow \eta\pi^0))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{330}/\Gamma$
VALUE (units $10^{-6}$ )				
<2.5	90	<sup>1</sup> AUBERT,BE	04 BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of charged and neutral $B$ mesons from $T(4S)$ decays.				

$\Gamma(a_0(980)^+ K^0 \times B(a_0(980)^+ \rightarrow \eta\pi^+))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{329}/\Gamma$
VALUE (units $10^{-6}$ )				
<3.9	90	<sup>1</sup> AUBERT,BE	04 BABR $e^+e^- \rightarrow T(4S)$	
<sup>1</sup> Assumes equal production of charged and neutral $B$ mesons from $T(4S)$ decays.				

$\Gamma(K^*(892)^0 \pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{331}/\Gamma$
VALUE (units $10^{-6}$ )				
<b>10.1 <math>\pm</math> 0.9 OUR AVERAGE</b>				
$10.8 \pm 0.6 \pm 1.2$ $-1.4$	<sup>1</sup> AUBERT	08A1	BABR $e^+e^- \rightarrow T(4S)$	
$9.67 \pm 0.64 \pm 0.81$ $-0.89$	<sup>1</sup> GARMASH	06	BELL $e^+e^- \rightarrow T(4S)$	

- • • We do not use the following data for averages, fits, limits, etc. • • •
- $13.5 \pm 1.2 \pm 0.8$   
 $-0.9$  <sup>1</sup> AUBERT,B 05N BABR Repl. by AUBERT 08A1
- $9.8 \pm 0.9 \pm 1.1$   
 $-1.2$  <sup>1</sup> GARMASH 05 BELL Repl. by GARMASH 06
- $15.5 \pm 1.8 \pm 1.5$   
 $-4.0$  <sup>1,2</sup> AUBERT,B 04P BABR Repl. by AUBERT,B 05N
- $19.4 \pm 4.2 \pm 4.1$   
 $-3.9 - 7.1$  <sup>3</sup> GARMASH 02 BELL Repl. by GARMASH 05
- <119 90 <sup>4</sup> ABE 00C SLD  $e^+e^- \rightarrow Z$
- < 16 90 <sup>1</sup> JESSOP 00 CLE2  $e^+e^- \rightarrow T(4S)$
- <390 90 <sup>5</sup> ADAM 96D DLPH  $e^+e^- \rightarrow Z$
- < 41 90 ASNER 96 CLE2 Repl. by JESSOP 00
- <480 90 <sup>5</sup> ABREU 95N DLPH Sup. by ADAM 96D
- <170 90 ALBRECHT 91B ARG  $e^+e^- \rightarrow T(4S)$
- <150 90 <sup>6</sup> AVERY 89B CLEO  $e^+e^- \rightarrow T(4S)$
- <260 90 AVERY 87 CLEO  $e^+e^- \rightarrow T(4S)$



$\Gamma(K_S^0(1430)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{342}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>45 <math>\pm \frac{9}{7}</math> OUR AVERAGE</b>	Error includes scale factor of 1.5.
$32.0 \pm 1.2 \pm^{+10.8}_{-6.0}$	<sup>1</sup> AUBERT 08A1 BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$51.6 \pm 1.7 \pm^{+7.0}_{-7.5}$	<sup>1</sup> GARMASH 06 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$44.4 \pm 2.2 \pm 5.3$	<sup>1,2</sup> AUBERT,B 05N BABR Repl. by AUBERT 08A1
$45.0 \pm 2.9 \pm^{+15.0}_{-10.7}$	<sup>1</sup> GARMASH 05 BELL Repl. by GARMASH 06

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> See erratum: AUBERT,BE 06A.

$\Gamma(K_S^*(1430)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{343}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>5.6 <math>\pm \frac{2.2}{1.5} \pm 0.1</math></b>	<sup>1,2</sup> AUBERT 08A1 BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
< 23	90 <sup>3</sup> AUBERT,B 05N BABR Repl. by AUBERT 08A1
< 6.9	90 <sup>4</sup> GARMASH 05 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< 680	90 ALBRECHT 91B ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> AUBERT 08A1 reports  $(1.85 \pm 0.41 \pm^{+0.61}_{-0.29}) \times 10^{-6}$  for  $B(B^+ \rightarrow K_S^*(1430)^0 \pi^+) \times B(K_S^*(1430)^0 \rightarrow K^+ \pi^-)$ . We compute  $B(B^+ \rightarrow K_S^*(1430)^0 \pi^+)$  using the PDG value  $B(K_S^*(1430)^0 \rightarrow K \pi) = (49.9 \pm 1.2) \times 10^{-2}$  and 2/3 for the  $K^+ \pi^-$  fraction. Our first error is their experiment's error and the second error is systematic error from using our best value.

<sup>3</sup> AUBERT,B 05N reports  $7.7 \times 10^{-6}$  at 90% CL for  $B(B^+ \rightarrow K_S^*(1430)^0 \pi^+) \times B(K_S^*(1430)^0 \rightarrow K^+ \pi^-)$ . We rescaled it using the PDG value  $B(K_S^*(1430)^0 \rightarrow K \pi) = 49.9\%$  and 2/3 for the  $K^+ \pi^-$  fraction.

<sup>4</sup> GARMASH 05 reports  $2.3 \times 10^{-6}$  at 90% CL for  $B(B^+ \rightarrow K_S^*(1430)^0 \pi^+) \times B(K_S^*(1430)^0 \rightarrow K^+ \pi^-)$ . We rescaled it using the PDG value  $B(K_S^*(1430)^0 \rightarrow K \pi) = 49.9\%$  and 2/3 for the  $K^+ \pi^-$  mode.

$\Gamma(K^*(1410)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{344}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>&lt;45</b>	90 <sup>1</sup> GARMASH 05 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> GARMASH 05 reports $2.0 \times 10^{-6}$ at 90% CL for $B(B^+ \rightarrow K^*(1410)^0 \pi^+) \times B(K^*(1410)^0 \rightarrow K^+ \pi^-)$ . We rescaled it using the PDG value $B(K^*(1410)^0 \rightarrow K \pi) = 6.6\%$ and 2/3 for the $K^+ \pi^-$ mode.	

$\Gamma(K^*(1680)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{345}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>&lt;12</b>	90 <sup>1</sup> GARMASH 05 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<15	90 <sup>2</sup> AUBERT,B 05N BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> GARMASH 05 reports $3.1 \times 10^{-6}$ at 90% CL for $B(B^+ \rightarrow K^*(1680)^0 \pi^+) \times B(K^*(1680)^0 \rightarrow K^+ \pi^-)$ . We rescaled it using the PDG value $B(K^*(1680)^0 \rightarrow K \pi) = 38.7\%$ and 2/3 for the $K^+ \pi^-$ mode.	
<sup>2</sup> AUBERT,B 05N reports $3.8 \times 10^{-6}$ at 90% CL for $B(B^+ \rightarrow K^*(1680)^0 \pi^+) \times B(K^*(1680)^0 \rightarrow K^+ \pi^-)$ . We rescaled it using the PDG value $B(K^*(1680)^0 \rightarrow K \pi) = 38.7\%$ and 2/3 for the $K^+ \pi^-$ fraction.	

$\Gamma(K^+ \pi^0 \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{346}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>16.2 <math>\pm 1.2 \pm 1.5</math></b>	<sup>1</sup> LEES 11i BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(f_0(980) K^+ \times B(f_0 \rightarrow \pi^0 \pi^0))/\Gamma_{\text{total}}$	$\Gamma_{347}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>2.8 <math>\pm 0.6 \pm 0.5</math></b>	<sup>1</sup> LEES 11i BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K^- \pi^+ \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{348}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>&lt;0.95</b>	90 <sup>1</sup> AUBERT 08BE BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<4.5	90 <sup>1</sup> GARMASH 04 BELL $e^+ e^- \rightarrow \Upsilon(4S)$
<1.8	90 <sup>2</sup> AUBERT 03M BABR Repl. by AUBERT 08BE
<7.0	90 <sup>3</sup> GARMASH 02 BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ ; charm and charmonium contributions are subtracted, otherwise no assumptions about intermediate resonances.

<sup>3</sup> Uses a reference decay mode  $B^+ \rightarrow \bar{D}^0 \pi^+$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  with  $B(B^+ \rightarrow \bar{D}^0 \pi^+) \cdot B(\bar{D}^0 \rightarrow K^+ \pi^-) = (20.3 \pm 2.0) \times 10^{-5}$ .

$\Gamma(K^- \pi^+ \pi^+ \text{ nonresonant})/\Gamma_{\text{total}}$	$\Gamma_{349}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>&lt;56</b>	90 BERGFELD 96B CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
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$\Gamma(K_1(1270)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{350}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT

<b>&lt;4.0 <math>\times 10^{-5}</math></b>	90 <sup>1</sup> AUBERT 10D BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K_1(1400)^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{351}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT

<b>&lt;3.9 <math>\times 10^{-5}</math></b>	90 <sup>1</sup> AUBERT 10D BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<2.6 $\times 10^{-3}$	90 ALBRECHT 91B ARG $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K^0 \pi^+ \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{352}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT

<b>&lt;66 <math>\times 10^{-6}</math></b>	90 <sup>1</sup> ECKHART 02 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K^0 \rho^+)/\Gamma_{\text{total}}$	$\Gamma_{353}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>8.0 <math>\pm \frac{1.4}{1.3} \pm 0.6</math></b>	AUBERT 07Z BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<48	90 ASNER 96 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(K^*(892)^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{354}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>75.3 <math>\pm 6.0 \pm 8.1</math></b>	<sup>1</sup> AUBERT,B 06U BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<1100	90 ALBRECHT 91E ARG $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K^*(892)^+ \rho^0)/\Gamma_{\text{total}}$	$\Gamma_{355}/\Gamma$
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT

<b>4.6 <math>\pm 1.0 \pm 0.4</math></b>	<sup>1</sup> DEL-AMO-SA..11D BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
< 6.1	90 <sup>1</sup> AUBERT,B 06G BABR Repl. by DEL-AMO-SANCHEZ 11D
$10.6 \pm \frac{3.0}{2.6} \pm 2.4$	<sup>1</sup> AUBERT 03V BABR Repl. by AUBERT,B 06G
< 74	90 <sup>2</sup> GODANG 02 CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
<900	90 ALBRECHT 91B ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to  $4.9 \times 10^{-5}$ .

$\Gamma(K^*(892)^+ f_0(980))/\Gamma_{\text{total}}$	$\Gamma_{356}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>4.2 <math>\pm 0.6 \pm 0.3</math></b>	<sup>1</sup> DEL-AMO-SA..11D BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
5.2 $\pm 1.2 \pm 0.5$	<sup>1</sup> AUBERT,B 06G BABR Repl. by DEL-AMO-SANCHEZ 11D
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(a_1^+ K^0)/\Gamma_{\text{total}}$	$\Gamma_{357}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>34.9 <math>\pm 5.0 \pm 4.4</math></b>	<sup>1,2</sup> AUBERT 08F BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	
<sup>2</sup> Assumes $a_1^\pm$ decays only to $3\pi$ and $B(a_1^\pm \rightarrow \pi^\pm \pi^\mp \pi^\pm) = 0.5$ .	

$\Gamma(b_1^+ K^0 \times B(b_1^+ \rightarrow \omega \pi^+))/\Gamma_{\text{total}}$	$\Gamma_{358}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>9.6 <math>\pm 1.7 \pm 0.9</math></b>	<sup>1</sup> AUBERT 08AG BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K^*(892)^0 \rho^+)/\Gamma_{\text{total}}$	$\Gamma_{359}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT

<b>9.2 <math>\pm 1.5</math> OUR AVERAGE</b>	
9.6 $\pm 1.7 \pm 1.5$	<sup>1</sup> AUBERT,B 06G BABR $e^+ e^- \rightarrow \Upsilon(4S)$
8.9 $\pm 1.7 \pm 1.2$	<sup>1</sup> ZHANG 05D BELL $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	

$\Gamma(K_1(1400)^+ \rho^0)/\Gamma_{\text{total}}$	$\Gamma_{360}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT

<b>&lt;7.8 <math>\times 10^{-4}</math></b>	90 ALBRECHT 91B ARG $e^+ e^- \rightarrow \Upsilon(4S)$
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$\Gamma(K_2^*(1430)^+ \rho^0)/\Gamma_{\text{total}}$	$\Gamma_{361}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT

<b>&lt;1.5 <math>\times 10^{-3}</math></b>	90 ALBRECHT 91B ARG $e^+ e^- \rightarrow \Upsilon(4S)$
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# Meson Particle Listings

$B^\pm$

$\Gamma(b_1^0 K^+ \times B(b_1^0 \rightarrow \omega \pi^0))/\Gamma_{\text{total}}$					$\Gamma_{362}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>9.1 \pm 1.7 \pm 1.0</math></b>	<sup>1</sup> AUBERT	07BI	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(b_1^+ K^{*0} \times B(b_1^+ \rightarrow \omega \pi^+))/\Gamma_{\text{total}}$					$\Gamma_{363}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 5.9 \times 10^{-6}</math></b>	90	<sup>1</sup> AUBERT	09AF	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(b_1^0 K^{*+} \times B(b_1^0 \rightarrow \omega \pi^0))/\Gamma_{\text{total}}$					$\Gamma_{364}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 6.7 \times 10^{-6}</math></b>	90	<sup>1</sup> AUBERT	09AF	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+ \bar{K}^0)/\Gamma_{\text{total}}$					$\Gamma_{365}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.31 \pm 0.17</math> OUR FIT</b>		Error includes scale factor of 1.2.			
<b><math>1.19 \pm 0.18</math> OUR AVERAGE</b>					
$1.11 \pm 0.19 \pm 0.05$		<sup>1</sup> DUH	13	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.61 \pm 0.44 \pm 0.09$		<sup>1</sup> AUBERT,BE	06c	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$1.22^{+0.32+0.13}_{-0.28-0.16}$		<sup>1</sup> LIN	07	BELL	Repl. by DUH 13
$1.0 \pm 0.4 \pm 0.1$		<sup>1</sup> ABE	05G	BELL	Repl. by LIN 07
$1.5 \pm 0.5 \pm 0.1$		<sup>1</sup> AUBERT,BE	05E	BABR	Repl. by AUBERT,BE 06C
$< 2.5$	90	<sup>1</sup> AUBERT	04M	BABR	Repl. by AUBERT,BE 05E
$< 3.3$	90	<sup>1</sup> CHAO	04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 3.3$	90	<sup>1</sup> BORNHEIM	03	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 2.0$	90	<sup>1</sup> CASEY	02	BELL	Repl. by CHAO 04
$< 5.0$	90	<sup>1</sup> ABE	01H	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 2.4$	90	<sup>1</sup> AUBERT	01E	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 5.1$	90	<sup>1</sup> CRONIN-HEN..00	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	
$< 21$	90	GODANG	98	CLE2	Repl. by CRONIN-HENNESSY 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+\bar{K}^0)/\Gamma(K^0\pi^+)$				$\Gamma_{365}/\Gamma_{305}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.055 \pm 0.007</math> OUR FIT</b>	Error includes scale factor of 1.2.			
<b><math>0.064 \pm 0.009 \pm 0.004</math></b>	AAIJ	13BS LHCB	$pp$ at 7 TeV	

$\Gamma(\bar{K}^0 K^+ \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{366}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 24 \times 10^{-6}</math></b>	90	<sup>1</sup> ECKHART	02	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+ K_S^0 \bar{K}_S^0)/\Gamma_{\text{total}}$	$\Gamma_{367}/\Gamma$			
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>10.8 \pm 0.6</math> OUR AVERAGE</b>				
$10.6 \pm 0.5 \pm 0.3$	<sup>1,2</sup> LEES	12o	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$13.4 \pm 1.9 \pm 1.5$	<sup>1</sup> GARMASH	04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$10.7 \pm 1.2 \pm 1.0$	<sup>1</sup> AUBERT,B	04v	BABR	Repl. by LEES 12o
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> All intermediate charmonium and charm resonances are removed, except of $\chi_{c0}$ .				

$\Gamma(f_0(980) K^+, f_0 \rightarrow K_S^0 \bar{K}_S^0)/\Gamma_{\text{total}}$					$\Gamma_{368}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>14.7 \pm 2.8 \pm 1.8</math></b>	<sup>1</sup> LEES	12O	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(f_0(1710) K^+, f_0 \rightarrow K_S^0 \bar{K}_S^0)/\Gamma_{\text{total}}$					$\Gamma_{369}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>0.40^{+0.40}_{-0.24} \pm 0.11</math></b>	<sup>1</sup> LEES	12O	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+ K_S^0 \bar{K}_S^0 \text{ nonresonant})/\Gamma_{\text{total}}$					$\Gamma_{370}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>19.8 \pm 3.7 \pm 2.5</math></b>	<sup>1</sup> LEES	12O	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_S^0 \bar{K}_S^0 \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{371}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 0.51</math></b>	90	<sup>1</sup> AUBERT	09J	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 3.2$	90	<sup>1</sup> GARMASH	04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+ K^- \pi^+)/\Gamma_{\text{total}}$					$\Gamma_{372}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>5.0 \pm 0.5 \pm 0.5</math></b>		<sup>1</sup> AUBERT	07BB	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 13$	90	<sup>1</sup> GARMASH	04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 6.3$	90	<sup>1,2</sup> AUBERT	03M	BABR	Repl. by AUBERT 07BB
$< 12$	90	<sup>3</sup> GARMASH	02	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Charm and charmonium contributions are subtracted, otherwise no assumptions about intermediate resonances.					
<sup>3</sup> Uses a reference decay mode $B^+ \rightarrow \bar{D}^0 \pi^+$ and $\bar{D}^0 \rightarrow K^+ \pi^-$ with $B(B^+ \rightarrow \bar{D}^0 \pi^+) \cdot B(\bar{D}^0 \rightarrow K^+ \pi^-) = (20.3 \pm 2.0) \times 10^{-5}$ .					

$\Gamma(K^+ K^- \pi^+ \text{ nonresonant})/\Gamma_{\text{total}}$					$\Gamma_{373}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 75</math></b>	90	BERGFELD	96B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

$\Gamma(K^+ \bar{K}^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{374}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 1.1</math></b>	90	<sup>1</sup> AUBERT	07AR	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 129$	90	ABBIENDI	00B	OPAL	$e^+ e^- \rightarrow Z$
$< 138$	90	<sup>2</sup> ABE	00C	SLD	$e^+ e^- \rightarrow Z$
$< 5.3$	90	<sup>1</sup> JESSOP	00	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> ABE 00C assumes $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$ and the $B$ fractions $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$ and $f_{B_s} = (10.5^{+1.8}_{-2.2})\%$ .					

$\Gamma(K^+ \bar{K}_S^0(1430)^0)/\Gamma_{\text{total}}$					$\Gamma_{375}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 2.2</math></b>	90	<sup>1</sup> AUBERT	07AR	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+ K^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{376}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 1.6 \times 10^{-7}</math></b>	90	<sup>1</sup> AUBERT	08BE	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 2.4 \times 10^{-6}$	90	<sup>1</sup> GARMASH	04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.3 \times 10^{-6}$	90	<sup>2</sup> AUBERT	03M	BABR	Repl. by AUBERT 08BE
$< 3.2 \times 10^{-6}$	90	<sup>3</sup> GARMASH	02	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Assumes equal production of $B^0$ and $B^+$ at the $\Upsilon(4S)$ ; charm and charmonium contributions are subtracted, otherwise no assumptions about intermediate resonances.					
<sup>3</sup> Uses a reference decay mode $B^+ \rightarrow \bar{D}^0 \pi^+$ and $\bar{D}^0 \rightarrow K^+ \pi^-$ with $B(B^+ \rightarrow \bar{D}^0 \pi^+) \cdot B(\bar{D}^0 \rightarrow K^+ \pi^-) = (20.3 \pm 2.0) \times 10^{-5}$ .					

$\Gamma(K^+ K^+ \pi^- \text{ nonresonant})/\Gamma_{\text{total}}$					$\Gamma_{377}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 87.9</math></b>	90	ABBIENDI	00B	OPAL	$e^+ e^- \rightarrow Z$

$\Gamma(f_2'(1525) K^+)/\Gamma_{\text{total}}$					$\Gamma_{378}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.8 \pm 0.5</math> OUR AVERAGE</b>		Error includes scale factor of 1.1.			
$1.56 \pm 0.36 \pm 0.30$		<sup>1,2</sup> LEES	12O	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$2.8 \pm 0.9^{+0.5}_{-0.4}$		<sup>1,3</sup> LEES	12O	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<8$	90	<sup>1,4</sup> GARMASH	05	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Measured in the $B^+ \rightarrow K^+ K^- K^+$ decay.					
<sup>3</sup> Measured in the $B^+ \rightarrow K^+ K_S^0 \bar{K}_S^0$ decay.					
<sup>4</sup> GARMASH 05 reports $B(B^+ \rightarrow f_2'(1525) K^+) \cdot B(f_2'(1525) \rightarrow K^+ K^-) < 4.9 \times 10^{-6}$ at 90% CL. We divide this result by our best value of $B(f_2'(1525) \rightarrow K \bar{K}) = 88.7 \times 10^{-2}$ multiplied by 2/3 to account for the $K^+ K^-$ fraction.					

$\Gamma(K^+ f_J(2220))/\Gamma_{\text{total}}$					$\Gamma_{379}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
not seen	<sup>1</sup> HUANG	03	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> No evidence is found for such decay and set a limit on $B(B^+ \rightarrow f_J(2220)) \times B(f_J(2220) \rightarrow \phi\phi) < 1.2 \times 10^{-6}$ at 90%CL where the $f_J(2220)$ is a possible glueball state.					

$\Gamma(K^{*+} \pi^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{380}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 11.8</math></b>	90	<sup>1</sup> AUBERT, B	06U	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(892)^+ K^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{381}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.91 \pm 0.29</math> OUR AVERAGE</b>				
$0.77^{+0.35}_{-0.30} \pm 0.12$		<sup>1</sup> GOH	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.2 \pm 0.5 \pm 0.1$		<sup>2</sup> AUBERT	09F BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<71	90	<sup>3</sup> GODANG	02 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Signal significance is 2.7 standard deviations. This measurement corresponds to an upper limit of $< 1.31 \times 10^{-6}$ at 90% CL. <sup>2</sup> Signal significance is 3.7 standard deviations. <sup>3</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to $4.8 \times 10^{-5}$ .				

 $\Gamma(K^+ K^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{382}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.1</b>	90	<sup>1</sup> AUBERT,B	06U BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(K^+ K^- K^+)/\Gamma_{\text{total}}$   $\Gamma_{383}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>34.0 \pm 1.4</math> OUR AVERAGE</b>				Error includes scale factor of 1.4.
$34.6 \pm 0.6 \pm 0.9$		<sup>1,2</sup> LEES	12o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$30.6 \pm 1.2 \pm 2.3$		<sup>1</sup> GARMASH	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$35.2 \pm 0.9 \pm 1.6$		<sup>1</sup> AUBERT	06o BABR	Repl. by LEES 12o
$32.8 \pm 1.8 \pm 2.8$		<sup>1</sup> GARMASH	04 BELL	Repl. by GARMASH 05
$29.6 \pm 2.1 \pm 1.6$		<sup>3</sup> AUBERT	03M BABR	Repl. by AUBERT 06o
$35.3 \pm 3.7 \pm 4.5$		<sup>4</sup> GARMASH	02 BELL	Repl. by GARMASH 04
<200	90	<sup>5</sup> ADAM	96D DLPH	$e^+ e^- \rightarrow Z$
<320	90	<sup>5</sup> ABREU	95N DLPH	Sup. by ADAM 96D
<350	90	ALBRECHT	91E ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>2</sup> All intermediate charmonium and charm resonances are removed, except of $\chi_{c0}$ . <sup>3</sup> Assumes equal production of $B^0$ and $B^+$ at the $\Upsilon(4S)$ ; charm and charmonium contributions are subtracted, otherwise no assumptions about intermediate resonances. <sup>4</sup> Uses a reference decay mode $B^+ \rightarrow \bar{D}^0 \pi^+$ and $\bar{D}^0 \rightarrow K^+ \pi^-$ with $B(B^+ \rightarrow \bar{D}^0 \pi^+) \cdot B(\bar{D}^0 \rightarrow K^+ \pi^-) = (20.3 \pm 2.0) \times 10^{-5}$ . <sup>5</sup> Assumes $B^0$ and $B^-$ production fractions of 0.39, and $B_S$ production fraction of 0.12.				

 $\Gamma(K^+ \phi)/\Gamma_{\text{total}}$   $\Gamma_{384}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>8.8^{+0.7}_{-0.6}</math> OUR AVERAGE</b>				Error includes scale factor of 1.1.
$9.2 \pm 0.4^{+0.7}_{-0.5}$		<sup>1</sup> LEES	12o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$7.6 \pm 1.3 \pm 0.6$		<sup>2</sup> ACOSTA	05J CDF	$p\bar{p}$ at 1.96 TeV
$9.60 \pm 0.92^{+1.05}_{-0.85}$		<sup>1</sup> GARMASH	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$5.5^{+2.1}_{-1.8} \pm 0.6$		<sup>1</sup> BRIERE	01 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.4 \pm 0.7 \pm 0.7$		<sup>1</sup> AUBERT	06o BABR	Repl. by LEES 12o
$10.0^{+0.9}_{-0.8} \pm 0.5$		<sup>1</sup> AUBERT	04A BABR	Repl. by AUBERT 06o
$9.4 \pm 1.1 \pm 0.7$		<sup>1</sup> CHEN	03B BELL	Repl. by GARMASH 05
$14.6^{+3.0}_{-2.8} \pm 2.0$		<sup>3</sup> GARMASH	02 BELL	Repl. by CHEN 03b
$7.7^{+1.6}_{-1.4} \pm 0.8$		<sup>1</sup> AUBERT	01D BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<144	90	<sup>4</sup> ABE	00c SLD	$e^+ e^- \rightarrow Z$
< 5	90	<sup>1</sup> BERGFELD	98 CLE2	
<280	90	<sup>5</sup> ADAM	96D DLPH	$e^+ e^- \rightarrow Z$
< 12	90	ASNER	96 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<440	90	<sup>6</sup> ABREU	95N DLPH	Sup. by ADAM 96D
<180	90	ALBRECHT	91B ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
< 90	90	<sup>7</sup> AVERY	89B CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
<210	90	AVERY	87 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>2</sup> Uses $B(B^+ \rightarrow J/\psi K^+) = (1.00 \pm 0.04) \times 10^{-3}$ and $B(J/\psi \rightarrow \mu^+ \mu^-) = 0.0588 \pm 0.0010$ . <sup>3</sup> Uses a reference decay mode $B^+ \rightarrow \bar{D}^0 \pi^+$ and $\bar{D}^0 \rightarrow K^+ \pi^-$ with $B(B^+ \rightarrow \bar{D}^0 \pi^+) \cdot B(\bar{D}^0 \rightarrow K^+ \pi^-) = (20.3 \pm 2.0) \times 10^{-5}$ . <sup>4</sup> ABE 00c assumes $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$ and the $B$ fractions $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$ and $f_{B_S} = (10.5^{+1.8}_{-2.2})\%$ . <sup>5</sup> ADAM 96D assumes $f_{B^0} = f_{B^-} = 0.39$ and $f_{B_S} = 0.12$ . <sup>6</sup> Assumes a $B^0$ , $B^-$ production fraction of 0.39 and a $B_S$ production fraction of 0.12. <sup>7</sup> AVERY 89B reports $< 8 \times 10^{-5}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0 \bar{B}^0$ . We rescale to 50%.				

 $\Gamma(f_0(980) K^+ \times B(f_0(980) \rightarrow K^+ K^-))/\Gamma_{\text{total}}$   $\Gamma_{385}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>9.4 \pm 1.6 \pm 2.8</math></b>		<sup>1</sup> LEES	12o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$6.5 \pm 2.5 \pm 1.6$		<sup>1</sup> AUBERT	06o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<2.9	90	<sup>1</sup> GARMASH	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(\phi_2(1320) K^+ \times B(\phi_2(1320) \rightarrow K^+ K^-))/\Gamma_{\text{total}}$   $\Gamma_{386}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> GARMASH	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(X_0(1550) K^+ \times B(X_0(1550) \rightarrow K^+ K^-))/\Gamma_{\text{total}}$   $\Gamma_{387}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.3 \pm 0.6 \pm 0.3</math></b>		<sup>1</sup> AUBERT	06o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(\phi(1680) K^+ \times B(\phi(1680) \rightarrow K^+ K^-))/\Gamma_{\text{total}}$   $\Gamma_{388}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.8 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> GARMASH	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(f_0(1710) K^+ \times B(f_0(1710) \rightarrow K^+ K^-))/\Gamma_{\text{total}}$   $\Gamma_{389}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.12 \pm 0.25 \pm 0.50</math></b>		<sup>1</sup> LEES	12o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.7 \pm 1.0 \pm 0.3$		<sup>1</sup> AUBERT	06o BABR	Repl. by LEES 12o
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(K^+ K^- K^+ \text{nonresonant})/\Gamma_{\text{total}}$   $\Gamma_{390}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>23.8^{+2.8}_{-5.0}</math> OUR AVERAGE</b>				
$22.8 \pm 2.7 \pm 7.6$		<sup>1</sup> LEES	12o BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$24.0 \pm 1.5^{+2.6}_{-6.0}$		<sup>1</sup> GARMASH	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$50.0 \pm 6.0 \pm 4.0$		<sup>1</sup> AUBERT	06o BABR	Repl. by LEES 12o
<38	90	BERGFELD	96B CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(K^*(892)^+ K^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{391}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>36.2 \pm 3.3 \pm 3.6</math></b>		<sup>1</sup> AUBERT,B	06U BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1600	90	ALBRECHT	91E ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(K^*(892)^+ \phi)/\Gamma_{\text{total}}$   $\Gamma_{392}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>10.0 \pm 2.0</math> OUR AVERAGE</b>				Error includes scale factor of 1.7.
$11.2 \pm 1.0 \pm 0.9$		<sup>1</sup> AUBERT	07BA BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$6.7^{+2.1}_{-1.9} \pm 0.7$		<sup>1</sup> CHEN	03B BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$12.7^{+2.2}_{-2.0} \pm 1.1$		<sup>1</sup> AUBERT	03V BABR	Repl. by AUBERT 07BA
$9.7^{+4.2}_{-3.4} \pm 1.7$		<sup>1</sup> AUBERT	01D BABR	Repl. by AUBERT 03v
< 22.5	90	<sup>1</sup> BRIERE	01 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
< 41	90	<sup>1</sup> BERGFELD	98 CLE2	
< 70	90	ASNER	96 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<1300	90	ALBRECHT	91B ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(\phi(K\pi)_0^{*+})/\Gamma_{\text{total}}$   $\Gamma_{393}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>8.3 \pm 1.4 \pm 0.8</math></b>		<sup>1</sup> AUBERT	08B1 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(\phi K_1(1270)^+)/\Gamma_{\text{total}}$   $\Gamma_{394}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>6.1 \pm 1.6 \pm 1.1</math></b>		<sup>1</sup> AUBERT	08B1 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(\phi K_1(1400)^+)/\Gamma_{\text{total}}$   $\Gamma_{395}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.2</b>	90	<sup>1</sup> AUBERT	08B1 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1100	90	ALBRECHT	91B ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

# Meson Particle Listings

$B^\pm$

$\Gamma(\phi K^*(1410)^+)/\Gamma_{\text{total}}$				$\Gamma_{396}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.3</b>	90	<sup>1</sup> AUBERT	08B1	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\phi K_0^*(1430)^+)/\Gamma_{\text{total}}$				$\Gamma_{397}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>7.0 \pm 1.3 \pm 0.9</math></b>		<sup>1</sup> AUBERT	08B1	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\phi K_2^*(1430)^+)/\Gamma_{\text{total}}$				$\Gamma_{398}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>8.4 \pm 1.8 \pm 1.0</math></b>		<sup>1</sup> AUBERT	08B1	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3400	90	ALBRECHT	91B	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\phi K_2^*(1770)^+)/\Gamma_{\text{total}}$				$\Gamma_{399}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;15.0</b>	90	<sup>1</sup> AUBERT	08B1	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\phi K_2^*(1820)^+)/\Gamma_{\text{total}}$				$\Gamma_{400}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;16.3</b>	90	<sup>1</sup> AUBERT	08B1	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(a_1^+ K^{*0})/\Gamma_{\text{total}}$				$\Gamma_{401}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.6</b>	90	<sup>1,2</sup> DEL-AMO-SA..101	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes  $B(a_1^\pm \rightarrow \pi^\pm \pi^\mp \pi^\pm) = 0.5$

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^+ \phi \phi)/\Gamma_{\text{total}}$				$\Gamma_{402}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.0 \pm 1.2</math> OUR AVERAGE</b>	Error includes scale factor of 2.3.			
$5.6 \pm 0.5 \pm 0.3$		<sup>1</sup> LEES	11A	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$2.6^{+1.1}_{-0.9} \pm 0.3$		<sup>1</sup> HUANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$7.5 \pm 1.0 \pm 0.7$		<sup>1</sup> AUBERT,BE	06H	BABR Repl. by LEES 11A

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$  and for a  $\phi\phi$  invariant mass below 2.85 GeV/ $c^2$ .

$\Gamma(\eta' \eta' K^+)/\Gamma_{\text{total}}$				$\Gamma_{403}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;25</b>	90	<sup>1</sup> AUBERT,B	06P	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\omega \phi K^+)/\Gamma_{\text{total}}$				$\Gamma_{404}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.9</b>	90	<sup>1</sup> LIU	09	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(1812) K^+ \times B(X \rightarrow \omega \phi))/\Gamma_{\text{total}}$				$\Gamma_{405}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.32</b>	90	<sup>1</sup> LIU	09	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{406}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.21 \pm 0.18</math> OUR AVERAGE</b>				
$4.22 \pm 0.14 \pm 0.16$		<sup>1</sup> AUBERT	09A0	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$4.25 \pm 0.31 \pm 0.24$		<sup>2</sup> NAKAO	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$3.76^{+0.89}_{-0.83} \pm 0.28$		<sup>2</sup> COAN	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.87 \pm 0.28 \pm 0.26$		<sup>3</sup> AUBERT,BE	04A	BABR Repl. by AUBERT 09A0
$3.83 \pm 0.62 \pm 0.22$		<sup>2</sup> AUBERT	02C	BABR Repl. by AUBERT,BE 04A
$5.7 \pm 3.1 \pm 1.1$		<sup>4</sup> AMMAR	93	CLE2 Repl. by COAN 00
< 55	90	<sup>5</sup> ALBRECHT	89G	ARG $e^+e^- \rightarrow \Upsilon(4S)$
< 55	90	<sup>5</sup> AVERY	89B	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
<180	90	AVERY	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> Uses the production ratio of charged and neutral B from  $\Upsilon(4S)$  decays  $R^{+0} = 1.006 \pm 0.048$ .

<sup>4</sup> AMMAR 93 observed  $4.1 \pm 2.3$  events above background.

<sup>5</sup> Assumes the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ .

$\Gamma(K_1(1270)^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{407}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.3 \pm 0.9 \pm 0.9</math></b>		<sup>1</sup> YANG	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 9.9	90	<sup>1</sup> NISHIDA	02	BELL Repl. by YANG 05
<730	90	<sup>2</sup> ALBRECHT	89G	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ALBRECHT 89G reports < 0.0066 assuming the  $\Upsilon(4S)$  decays 45% to  $B^0 \bar{B}^0$ . We rescale to 50%.

$\Gamma(\eta K^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{408}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>7.9 \pm 0.9</math> OUR AVERAGE</b>				
$7.7 \pm 1.0 \pm 0.4$		<sup>1,2</sup> AUBERT	09	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$8.4 \pm 1.5^{+1.2}_{-0.9}$		<sup>2,3</sup> NISHIDA	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1,2</sup> AUBERT,B 06M BABR Repl. by AUBERT 09

<sup>1</sup>  $m_{\eta K} < 3.25$  GeV/ $c^2$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup>  $m_{\eta K} < 2.4$  GeV/ $c^2$

$\Gamma(\eta' K^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{409}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.9 \pm 1.0_{-0.9}</math> OUR AVERAGE</b>				
$3.6 \pm 1.2 \pm 0.4$		<sup>1,2</sup> WEDD	10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$1.9^{+1.5}_{-1.2} \pm 0.1$		<sup>1,3</sup> AUBERT,B	06M	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup>  $m_{\eta' K} < 3.4$  GeV/ $c^2$ .

<sup>3</sup> Set the upper limit of  $4.2 \times 10^{-6}$  at 90% CL with  $m_{\eta' K} < 3.25$  GeV/ $c^2$ .

$\Gamma(\phi K^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{410}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.7 \pm 0.4</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.			
$2.48 \pm 0.30 \pm 0.24$		<sup>1</sup> SAHOO	11A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$3.5 \pm 0.6 \pm 0.4$		<sup>1</sup> AUBERT	07Q	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup> DRUTSKOY 04 BELL Repl. by SAHOO 11A

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at  $\Upsilon(4S)$ .

$\Gamma(K^+ \pi^- \pi^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{411}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.76 \pm 0.22</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.			
$2.95 \pm 0.13 \pm 0.20$		<sup>1,2</sup> AUBERT	07R	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$2.50 \pm 0.18 \pm 0.22$		<sup>2,3</sup> YANG	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>2,4</sup> NISHIDA 02 BELL Repl. by YANG 05

<sup>1</sup>  $M_{K\pi\pi} < 1.8$  GeV/ $c^2$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup>  $M_{K\pi\pi} < 2.0$  GeV/ $c^2$ .

<sup>4</sup>  $M_{K\pi\pi} < 2.4$  GeV/ $c^2$ .

$\Gamma(K^*(892)^0 \pi^+ \gamma)/\Gamma_{\text{total}}$				$\Gamma_{412}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>(2.0 \pm 0.7 \pm 0.2) \times 10^{-5}</math></b>		<sup>1,2</sup> NISHIDA	02	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup>  $M_{K\pi\pi} < 2.4$  GeV/ $c^2$ .

$\Gamma(K^+ \rho^0 \gamma)/\Gamma_{\text{total}}$				$\Gamma_{413}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.0 <math>\times 10^{-5}</math></b>	90	<sup>1,2</sup> NISHIDA	02	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup>  $M_{K\pi\pi} < 2.4$  GeV/ $c^2$ .

$\Gamma(K^+ \pi^- \pi^+ \gamma \text{ nonresonant})/\Gamma_{\text{total}}$				$\Gamma_{414}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;9.2 <math>\times 10^{-6}</math></b>	90	<sup>1,2</sup> NISHIDA	02	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup>  $M_{K\pi\pi} < 2.4$  GeV/ $c^2$ .

$\Gamma(K^0 \pi^+ \pi^0 \gamma)/\Gamma_{\text{total}}$				$\Gamma_{415}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.56 \pm 0.42 \pm 0.31</math></b>		<sup>1,2</sup> AUBERT	07R	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>  $M_{K\pi\pi} < 1.8$  GeV/ $c^2$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

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## Meson Particle Listings

 $B^\pm$  $\Gamma(K_1(1400)^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{416}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 1.5	90	<sup>1</sup> YANG 05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 5.0	90	<sup>1</sup> NISHIDA 02	BELL	Repl. by YANG 05
<220	90	<sup>2</sup> ALBRECHT 89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 89G reports < 0.0020 assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(K_2^*(1430)^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{417}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.45 ± 0.40 ± 0.15</b>		<sup>1</sup> AUBERT,B 04U	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<140	90	<sup>2</sup> ALBRECHT 89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 89G reports < 0.0013 assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(K^*(1680)^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{418}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.0019	90	<sup>1</sup> ALBRECHT 89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 89G reports < 0.0017 assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(K_3^*(1780)^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{419}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 39	90	<sup>1,2</sup> NISHIDA 05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5500	90	<sup>3</sup> ALBRECHT 89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $B(K_3^*(1780) \rightarrow \eta K) = 0.11^{+0.05}_{-0.04}$ .  
<sup>3</sup> ALBRECHT 89G reports < 0.005 assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(K_4^*(2045)^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{420}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.0099	90	<sup>1</sup> ALBRECHT 89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 89G reports < 0.0090 assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(\rho^+\gamma)/\Gamma_{\text{total}}$   $\Gamma_{421}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.98 ± 0.25 OUR AVERAGE</b>				
$1.20^{+0.42}_{-0.37} \pm 0.20$		<sup>1</sup> AUBERT 08BH	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.87^{+0.29+0.09}_{-0.27-0.11}$		<sup>1</sup> TANIGUCHI 08	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.10^{+0.37}_{-0.33} \pm 0.09$		<sup>1</sup> AUBERT 07L	BABR	Repl. by AUBERT 08BH
$0.55^{+0.42+0.09}_{-0.36-0.08}$		<sup>1</sup> MOHAPATRA 06	BELL	Repl. by TANIGUCHI 08
$0.9^{+0.6}_{-0.5} \pm 0.1$	90	<sup>1</sup> AUBERT 05	BABR	Repl. by AUBERT 07L
< 2.2	90	<sup>1</sup> MOHAPATRA 05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 2.1	90	<sup>1</sup> AUBERT 04C	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<13	90	<sup>1,2</sup> COAN 90	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at  $\Upsilon(4S)$ .  
<sup>2</sup> No evidence for a nonresonant  $K\pi\gamma$  contamination was seen; the central value assumes no contamination.

 $\Gamma(\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{422}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>5.5 ± 0.4 OUR AVERAGE</b>				Error includes scale factor of 1.2.
$5.86 \pm 0.26 \pm 0.38$		<sup>1</sup> DUH 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$5.02 \pm 0.46 \pm 0.29$		<sup>1</sup> AUBERT 07Bc	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$4.6^{+1.8+0.6}_{-1.6-0.7}$		<sup>1</sup> BORNHEIM 03	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$6.5 \pm 0.4 \pm 0.4$		<sup>1</sup> LIN 07A	BELL	Repl. by DUH 13
$5.8 \pm 0.6 \pm 0.4$		<sup>1</sup> AUBERT 05L	BABR	Repl. by AUBERT 07Bc
$5.0 \pm 1.2 \pm 0.5$		<sup>1</sup> CHAO 04	BELL	Repl. by LIN 07A
$5.5^{+1.0}_{-1.9} \pm 0.6$		<sup>1</sup> AUBERT 03L	BABR	Repl. by AUBERT 05L
$7.4^{+2.3}_{-2.2} \pm 0.9$		<sup>1</sup> CASEY 02	BELL	Repl. by CHAO 04
< 13.4	90	<sup>1</sup> ABE 01H	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 9.6	90	<sup>1</sup> AUBERT 01E	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 12.7	90	<sup>1</sup> CRONIN-HEN.00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 20	90	GODANG 98	CLE2	Repl. by CRONIN-HENNESSY 00
< 17	90	ASNER 96	CLE2	Repl. by GODANG 98
< 240	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<2300	90	<sup>2</sup> BEBEK 87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> BEBEK 87 assume the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ .

 $\Gamma(\pi^+\pi^0)/\Gamma(K^0\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.285 ± 0.02 ± 0.02</b>	LIN 07A	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma(\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{423}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>15.2 ± 0.6 ± 1.3 ± 1.2</b>		<sup>1</sup> AUBERT 09L	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$16.2 \pm 1.2 \pm 0.9$		<sup>1</sup> AUBERT,B 05G	BABR	Repl. by AUBERT 09L
$10.9 \pm 3.3 \pm 1.6$		<sup>1</sup> AUBERT 03M	BABR	Repl. by AUBERT 05G
<130	90	<sup>2</sup> ADAM 96D	DLPH	$e^+e^- \rightarrow Z$
<220	90	<sup>3</sup> ABREU 95N	DLPH	Sup. by ADAM 96D
<450	90	<sup>4</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<190	90	<sup>5</sup> BORTOLETTO 89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ ; charm and charmonium contributions are subtracted, otherwise no assumptions about intermediate resonances.  
<sup>2</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ .  
<sup>3</sup> Assumes a  $B^0$ ,  $B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12.  
<sup>4</sup> ALBRECHT 90B limit assumes equal production of  $B^0\bar{B}^0$  and  $B^+\pi^-$  at  $\Upsilon(4S)$ .  
<sup>5</sup> BORTOLETTO 89 reports <  $1.7 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(\rho^0\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{424}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>8.3 ± 1.2 OUR AVERAGE</b>				
$8.1 \pm 0.7^{+1.3}_{-1.6}$		<sup>1</sup> AUBERT 09L	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$8.0^{+2.3}_{-2.0} \pm 0.7$		<sup>1</sup> GORDON 02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$10.4^{+3.4}_{-3.4} \pm 2.1$		<sup>1</sup> JESSOP 00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.8 \pm 1.0^{+0.6}_{-0.9}$		<sup>1</sup> AUBERT,B 05G	BABR	Repl. by AUBERT 09L
$9.5 \pm 1.1 \pm 0.9$		<sup>1</sup> AUBERT 04Z	BABR	Repl. by AUBERT 05G
< 83	90	<sup>2</sup> ABE 00C	SLD	$e^+e^- \rightarrow Z$
<160	90	<sup>3</sup> ADAM 96D	DLPH	$e^+e^- \rightarrow Z$
< 43	90	ASNER 96	CLE2	Repl. by JESSOP 00
<260	90	<sup>4</sup> ABREU 95N	DLPH	Sup. by ADAM 96D
<150	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<170	90	<sup>5</sup> BORTOLETTO 89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<230	90	<sup>5</sup> BEBEK 87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<600	90	GILES 84	CLEO	Repl. by BEBEK 87

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ABE 00C assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$  and  $f_{B_s} = (10.5^{+1.8}_{-2.2})\%$ .  
<sup>3</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ .  
<sup>4</sup> Assumes a  $B^0$ ,  $B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12.  
<sup>5</sup> Papers assume the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $[\Gamma(K^*(892)^0\pi^+) + \Gamma(\rho^0\pi^+)]/\Gamma_{\text{total}}$   $(\Gamma_{331} + \Gamma_{424})/\Gamma$ 

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>170 ± 120 ± 80 ± 20</b>	<sup>1</sup> ADAM 96D	DLPH	$e^+e^- \rightarrow Z$

<sup>1</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ .

 $\Gamma(\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{425}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 1.5	90	<sup>1</sup> AUBERT 09L	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 3.0	90	<sup>1</sup> AUBERT,B 05G	BABR	Repl. by AUBERT 09L
<140	90	<sup>2</sup> BORTOLETTO 89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> BORTOLETTO 89 reports <  $1.2 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

 $\Gamma(\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{426}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.60 ± 0.67 ± 0.02 ± 0.44 ± 0.06</b>		<sup>1,2</sup> AUBERT 09L	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.10 \pm 1.28^{+0.04}_{-0.14}$		<sup>2,3</sup> AUBERT,B 05G	BABR	Repl. by AUBERT 09L
<240	90	<sup>4</sup> BORTOLETTO 89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 09L reports  $[\Gamma(B^+ \rightarrow \pi^+\pi^0(1270))/\Gamma_{\text{total}}] \times [B(\pi^0(1270) \rightarrow \pi^+\pi^-)] = (0.9 \pm 0.2 \pm 0.1^{+0.3}_{-0.1}) \times 10^{-6}$  which we divide by our best value  $B(\pi^0(1270) \rightarrow \pi^+\pi^-) = (56.2^{+1.9}_{-0.6}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> AUBERT,B 05G reports  $[\Gamma(B^+ \rightarrow \pi^+\pi^0(1270))/\Gamma_{\text{total}}] \times [B(\pi^0(1270) \rightarrow \pi^+\pi^-)] = (2.3 \pm 0.6 \pm 0.4) \times 10^{-6}$  which we divide by our best value  $B(\pi^0(1270) \rightarrow \pi^+\pi^-) = (56.2^{+1.9}_{-0.6}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> BORTOLETTO 89 reports <  $2.1 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

# Meson Particle Listings

$B^\pm$

$\Gamma(\rho(1450)^0\pi^+, \rho^0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{427}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

$1.4 \pm 0.4^{+0.5}_{-0.8}$		<sup>1</sup> AUBERT	09L	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<2.3	90	<sup>1</sup> AUBERT,B	05G	BABR	Repl. by AUBERT 09L
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(f_0(1370)\pi^+, f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{428}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<4.0	90	<sup>1</sup> AUBERT	09L	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<3.0	90	<sup>1</sup> AUBERT,B	05G	BABR	Repl. by AUBERT 09L
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(f_0(500)\pi^+, f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{429}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<4.1	90	<sup>1</sup> AUBERT,B	05G	BABR	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\pi^+\pi^-\pi^+\pi^0 \text{ nonresonant})/\Gamma_{\text{total}}$					$\Gamma_{430}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

$5.3 \pm 0.7^{+1.3}_{-0.8}$		<sup>1</sup> AUBERT	09L	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 4.6	90	<sup>1</sup> AUBERT,B	05G	BABR	Repl. by AUBERT 09L
<41	90	BERGFELD	96B	CLE2	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\pi^+\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{431}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	

<8.9 $\times 10^{-4}$	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+\pi^-$ at $\gamma(4S)$ .					

$\Gamma(\rho^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{432}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>10.9 <math>\pm 1.4</math> OUR AVERAGE</b>					
10.2 $\pm 1.4 \pm 0.9$		<sup>1</sup> AUBERT	07x	BABR	$e^+e^- \rightarrow \gamma(4S)$
13.2 $\pm 2.3^{+1.4}_{-1.9}$		<sup>1</sup> ZHANG	05A	BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
10.9 $\pm 1.9 \pm 1.9$		<sup>1</sup> AUBERT	04z	BABR	Repl. by AUBERT 07x
< 43	90	<sup>1,2</sup> JESSOP	00	CLE2	$e^+e^- \rightarrow \gamma(4S)$
< 77	90	ASNER	96	CLE2	Repl. by JESSOP 00
<550	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> Assumes no nonresonant contributions of $B^+ \rightarrow \pi^+\pi^0\pi^0$ .					

$\Gamma(\pi^+\pi^-\pi^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{433}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	

<4.0 $\times 10^{-3}$	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+\pi^-$ at $\gamma(4S)$ .					

$\Gamma(\rho^+\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{434}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>24.0 <math>\pm 1.9</math> OUR AVERAGE</b>					
23.7 $\pm 1.4 \pm 1.4$		<sup>1</sup> AUBERT	09G	BABR	$e^+e^- \rightarrow \gamma(4S)$
31.7 $\pm 7.1^{+3.8}_{-6.7}$		<sup>1,2</sup> ZHANG	03B	BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
16.8 $\pm 2.2 \pm 2.3$		<sup>1</sup> AUBERT,BE	06G	BABR	Repl. by AUBERT 09G
22.5 $\pm 5.7^{+3.8}_{-5.4} \pm 5.8$		<sup>1</sup> AUBERT	03v	BABR	Repl. by AUBERT,BE 06G
< 1000	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> The systematic error includes the error associated with the helicity-mix uncertainty.					

$\Gamma(\rho^+f_0(980), f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{435}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<2.0	90	<sup>1</sup> AUBERT	09G	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.9	90	<sup>1</sup> AUBERT,BE	06G	BABR	Repl. by AUBERT 09G
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(a_1(1260)^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{436}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

$26.4 \pm 5.4 \pm 4.1$		<sup>1,2</sup> AUBERT	07BL	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1700	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> Assumes $a_1^+$ decays only to $3\pi$ and $B(a_1^+ \rightarrow \pi^\pm\pi^\mp\pi^+) = 0.5$ .					

$\Gamma(a_1(1260)^0\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{437}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>20.4 <math>\pm 4.7 \pm 3.4</math></b>		<sup>1,2</sup> AUBERT	07BL	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<900	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> Assumes $a_1^0$ decays only to $3\pi$ and $B(a_1^+ \rightarrow \pi^\pm\pi^\mp\pi^0) = 1.0$ .					

$\Gamma(\omega\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{438}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>6.9 <math>\pm 0.5</math> OUR AVERAGE</b>					
6.7 $\pm 0.5 \pm 0.4$		<sup>1</sup> AUBERT	07AE	BABR	$e^+e^- \rightarrow \gamma(4S)$
6.9 $\pm 0.6 \pm 0.5$		<sup>1</sup> JEN	06	BELL	$e^+e^- \rightarrow \gamma(4S)$
11.3 $\pm 3.3^{+1.4}_{-2.9} \pm 1.4$		<sup>1</sup> JESSOP	00	CLE2	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
6.1 $\pm 0.7 \pm 0.4$		<sup>1</sup> AUBERT,B	06E	BABR	Repl. by AUBERT 07AE
5.5 $\pm 0.9 \pm 0.5$		<sup>1</sup> AUBERT	04H	BABR	Repl. by AUBERT,B 06E
5.7 $\pm 1.4^{+1.3}_{-1.3} \pm 0.6$		<sup>1</sup> WANG	04A	BELL	Repl. by JEN 06
4.2 $\pm 2.0^{+1.8}_{-1.8} \pm 0.5$		<sup>1</sup> LU	02	BELL	Repl. by WANG 04A
6.6 $\pm 2.1^{+1.8}_{-1.8} \pm 0.7$		<sup>1</sup> AUBERT	01G	BABR	Repl. by AUBERT 04H
< 23	90	<sup>1</sup> BERGFELD	98	CLE2	Repl. by JESSOP 00
<400	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\omega\rho^+)/\Gamma_{\text{total}}$					$\Gamma_{439}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>15.9 <math>\pm 1.6 \pm 1.4</math></b>		<sup>1</sup> AUBERT	09H	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
10.6 $\pm 2.1^{+1.6}_{-1.0}$		<sup>1</sup> AUBERT,B	06T	BABR	Repl. by AUBERT 09H
12.6 $\pm 3.7^{+3.3}_{-3.3} \pm 1.6$		<sup>1</sup> AUBERT	05o	BABR	Repl. by AUBERT,B 06T
<61	90	<sup>1</sup> BERGFELD	98	CLE2	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\eta\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{440}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>4.02 <math>\pm 0.27</math> OUR AVERAGE</b>					
4.07 $\pm 0.26 \pm 0.21$		<sup>1</sup> HOI	12	BELL	$e^+e^- \rightarrow \gamma(4S)$
4.00 $\pm 0.40 \pm 0.24$		<sup>1</sup> AUBERT	09AV	BABR	$e^+e^- \rightarrow \gamma(4S)$
1.2 $\pm 2.8^{+2.8}_{-1.2}$		<sup>1</sup> RICHICHI	00	CLE2	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
5.0 $\pm 0.5 \pm 0.3$		<sup>1</sup> AUBERT	07AE	BABR	Repl. by AUBERT 09AV
4.2 $\pm 0.4 \pm 0.2$		<sup>1</sup> CHANG	07B	BELL	Repl. by HOI 12
5.1 $\pm 0.6 \pm 0.3$		<sup>1</sup> AUBERT,B	05K	BABR	Repl. by AUBERT 07AE
4.8 $\pm 0.7 \pm 0.3$		<sup>1</sup> CHANG	05A	BELL	Repl. by CHANG 07B
5.3 $\pm 1.0 \pm 0.3$		<sup>1</sup> AUBERT	04H	BABR	Repl. by AUBERT,B 05K
< 15	90	BEHRENS	98	CLE2	Repl. by RICHICHI 00
<700	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\eta\rho^+)/\Gamma_{\text{total}}$					$\Gamma_{441}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>7.0 <math>\pm 2.9</math> OUR AVERAGE</b>		Error includes scale factor of 2.8.			
9.9 $\pm 1.2 \pm 0.8$		<sup>1</sup> AUBERT	08AH	BABR	$e^+e^- \rightarrow \gamma(4S)$
4.1 $\pm 1.4^{+1.4}_{-1.3} \pm 0.4$		<sup>1</sup> WANG	07B	BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
8.4 $\pm 1.9 \pm 1.1$		<sup>1</sup> AUBERT,B	05K	BABR	Repl. by AUBERT 08AH
<14	90	<sup>1</sup> AUBERT,B	04D	BABR	Repl. by AUBERT,B 05K
<15	90	<sup>1</sup> RICHICHI	00	CLE2	$e^+e^- \rightarrow \gamma(4S)$
<32	90	BEHRENS	98	CLE2	Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(\eta'\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{442}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<b>2.7 <math>\pm 0.9</math> OUR AVERAGE</b>		Error includes scale factor of 1.9.			
3.5 $\pm 0.6 \pm 0.2$		<sup>1</sup> AUBERT	09AV	BABR	$e^+e^- \rightarrow \gamma(4S)$
1.76 $\pm 0.67^{+0.15}_{-0.62} \pm 0.14$		<sup>1</sup> SCHUEMANN	06	BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3.9 $\pm 0.7 \pm 0.3$		<sup>1</sup> AUBERT	07AE	BABR	Repl. by AUBERT 09AV
4.0 $\pm 0.8 \pm 0.4$		<sup>1</sup> AUBERT,B	05K	BABR	Repl. by AUBERT 07AE
< 4.5	90	<sup>1</sup> AUBERT	04H	BABR	Repl. by AUBERT,B 05K
< 7.0	90	<sup>1</sup> ABE	01M	BELL	$e^+e^- \rightarrow \gamma(4S)$
<12	90	<sup>1</sup> AUBERT	01G	BABR	$e^+e^- \rightarrow \gamma(4S)$
<12	90	<sup>1</sup> RICHICHI	00	CLE2	$e^+e^- \rightarrow \gamma(4S)$
<31	90	BEHRENS	98	CLE2	Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					



$\Gamma(\eta'\rho^+)/\Gamma_{\text{total}}$				$\Gamma_{443}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$9.7^{+1.9}_{-1.8} \pm 1.1$		<sup>1</sup> DEL-AMO-SA...10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.7^{+3.1+2.3}_{-2.8-1.3}$		<sup>1</sup> AUBERT	07E BABR	Repl. by DEL-AMO-SANCHEZ 10A
< 5.8	90	<sup>1</sup> SCHUEMANN	07 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<22	90	<sup>1</sup> AUBERT,B	04D BABR	Repl. by AUBERT 07E
<33	90	<sup>1</sup> RICHICHI	00 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
<47	90	BEHRENS	98 CLE2	Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\phi\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{444}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>1.5</b>	90	<sup>1</sup> AAIJ	14A LHCB	$pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 3.3	90	<sup>2</sup> KIM	12A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 2.4	90	<sup>2</sup> AUBERT,B	06C BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 4.1	90	<sup>2</sup> AUBERT	04A BABR	Repl. by AUBERT,B 06C
< 14	90	<sup>2</sup> AUBERT	01D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<1530	90	<sup>3</sup> ABE	00C SLD	$e^+e^- \rightarrow Z$
< 50	90	<sup>2</sup> BERGFELD	98 CLE2	
<sup>1</sup> Measures $B(B^+ \rightarrow \phi\pi^+)/B(B^+ \rightarrow \phi K^+) < 0.018$ at 90% C.L. and assumes $B(B^+ \rightarrow \phi K^+) = (8.8^{+0.7}_{-0.6}) \times 10^{-6}$ .				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>3</sup> ABE 00C assumes $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$ and the $B$ fractions $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$ and $f_{B_s} = (10.5^{+1.8}_{-2.2})\%$ .				

$\Gamma(\phi\rho^+)/\Gamma_{\text{total}}$				$\Gamma_{445}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>3.0</b>	90	<sup>1</sup> AUBERT	08BK BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<16		<sup>1</sup> BERGFELD	98 CLE2	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(a_0(980)^0\pi^+, a_0^0 \rightarrow \eta\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{446}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>5.8</b>	90	<sup>1</sup> AUBERT,BE	04 BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of charged and neutral $B$ mesons from $\Upsilon(4S)$ decays.				

$\Gamma(a_0(980)^+\pi^0, a_0^+ \rightarrow \eta\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{447}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>1.4</b>	90	<sup>1</sup> AUBERT	08A BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\pi^+\pi^+\pi^-\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{448}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>8.6 \times 10^{-4}</math></b>	90	<sup>1</sup> ALBRECHT	90B ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .				

$\Gamma(\rho^0 a_1(1260)^+)/\Gamma_{\text{total}}$				$\Gamma_{449}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>6.2 \times 10^{-4}</math></b>	90	<sup>1</sup> BORTOLETTO89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< $6.0 \times 10^{-4}$	90	<sup>2</sup> ALBRECHT	90B ARG	$e^+e^- \rightarrow \Upsilon(4S)$
< $3.2 \times 10^{-3}$	90	<sup>1</sup> BEBEK	87 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> BORTOLETTO 89 reports $< 5.4 \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				
<sup>2</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .				

$\Gamma(\rho^0 a_2(1320)^+)/\Gamma_{\text{total}}$				$\Gamma_{450}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>7.2 \times 10^{-4}</math></b>	90	<sup>1</sup> BORTOLETTO89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< $2.6 \times 10^{-3}$	90	<sup>2</sup> BEBEK	87 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> BORTOLETTO 89 reports $< 6.3 \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				
<sup>2</sup> BEBEK 87 reports $< 2.3 \times 10^{-3}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				

$\Gamma(b_1^0\pi^+, b_1^0 \rightarrow \omega\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{451}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
<b><math>6.7 \pm 1.7 \pm 1.0</math></b>		<sup>1</sup> AUBERT	07BI BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(b_1^+\pi^0, b_1^+ \rightarrow \omega\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{452}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>3.3</b>	90	<sup>1</sup> AUBERT	08AG BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\pi^+\pi^+\pi^-\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{453}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>6.3 \times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT	90B ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .				

$\Gamma(b_1^+\rho^0, b_1^+ \rightarrow \omega\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{454}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>5.2 \times 10^{-6}</math></b>	90	<sup>1</sup> AUBERT	09AF BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(b_1^0\rho^+, b_1^0 \rightarrow \omega\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{456}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>3.3 \times 10^{-6}</math></b>	90	<sup>1</sup> AUBERT	09AF BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(a_1(1260)^0 a_1(1260)^0)/\Gamma_{\text{total}}$				$\Gamma_{455}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< <b><math>1.3 \times 10^{-2}</math></b>	90	<sup>1</sup> ALBRECHT	90B ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .				

$\Gamma(h^+\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{457}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
$16^{+6}_{-5} \pm 3.6$		GODANG	98 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$h^+ = K^+$ or $\pi^+$				

$\Gamma(\omega h^+)/\Gamma_{\text{total}}$				$\Gamma_{458}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
<b><math>13.8^{+2.7}_{-2.4}</math> OUR AVERAGE</b>				
$13.4^{+3.3}_{-2.9} \pm 1.1$		<sup>1</sup> LU	02 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$14.3^{+3.6}_{-3.2} \pm 2.0$		<sup>1</sup> JESSOP	00 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$25^{+8}_{-7} \pm 3$		<sup>1</sup> BERGFELD	98 CLE2	Repl. by JESSOP 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(h^+X^0(\text{Familon}))/\Gamma_{\text{total}}$				$\Gamma_{459}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>49</b>	90	<sup>1</sup> AMMAR	01B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> AMMAR 01B searched for the two-body decay of the $B$ meson to a massless neutral feebly-interacting particle $X^0$ such as the familon, the Nambu-Goldstone boson associated with a spontaneously broken global family symmetry.				

$\Gamma(\rho\bar{p}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{460}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.62 \pm 0.20</math> OUR AVERAGE</b>				
$1.60^{+0.22}_{-0.19} \pm 0.12$		<sup>1,2,3</sup> WEI	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.69 \pm 0.29 \pm 0.26$		<sup>1</sup> AUBERT	07Av BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.07 \pm 0.11 \pm 0.11$		<sup>4</sup> AAIJ	14AF LHCB	$pp$ at 7, 8 TeV
$3.06^{+0.73}_{-0.62} \pm 0.37$		<sup>1,3</sup> WANG	04 BELL	Repl. by WEI 08
< 3.7	90	<sup>1,2</sup> ABE	02K BELL	Repl. by WANG 04
<500	90	<sup>5</sup> ABREU	95N DLPH	Repl. by ADAM 96D
<160	90	<sup>6</sup> BEBEK	89 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
$570 \pm 150 \pm 210$		<sup>7</sup> ALBRECHT	88F ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> Explicitly vetoes resonant production of $p\bar{p}$ from Charmonium states.				
<sup>3</sup> Also provides results with $m_{\rho\bar{p}} < 2.85 \text{ GeV}/c^2$ and angular asymmetry of $p\bar{p}$ system.				
<sup>4</sup> Requires $m_{p\bar{p}} < 2.85 \text{ GeV}/c^2$ .				
<sup>5</sup> Assumes a $B^0, B^-$ production fraction of 0.39 and a $B_s$ production fraction of 0.12.				
<sup>6</sup> BEBEK 89 reports $< 1.4 \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				
<sup>7</sup> ALBRECHT 88F reports $(5.2 \pm 1.4 \pm 1.9) \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 45% to $B^0\bar{B}^0$ . We rescale to 50%.				

$\Gamma(\rho\bar{p}\pi^+\text{nonresonant})/\Gamma_{\text{total}}$				$\Gamma_{461}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< <b>53</b>	90	BERGFELD	96B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\rho\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{462}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< $5.2 \times 10^{-4}$	90	<sup>1</sup> ALBRECHT	88F ARG	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 88F reports $< 4.7 \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 45% to $B^0\bar{B}^0$ . We rescale to 50%.				

# Meson Particle Listings

$B^\pm$

$(p\overline{p}K^+)/\Gamma_{\text{total}}$				$\Gamma_{463}/\Gamma$			
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT			
<b><math>5.9 \pm 0.5</math></b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.5.					
$5.54^{+0.27}_{-0.25} \pm 0.36$		1,2,3 WEI	08	BELL	$e^+e^- \rightarrow \gamma(4S)$		
$6.7 \pm 0.5 \pm 0.4$		1,3 AUBERT,B	05L	BABR	$e^+e^- \rightarrow \gamma(4S)$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
$4.59^{+0.38}_{-0.34} \pm 0.50$		1,2,3 WANG	05A	BELL	Repl. by WEI 08		
$5.66^{+0.67}_{-0.57} \pm 0.62$		1,2,3 WANG	04	BELL	Repl. by WANG 05A		
$4.3^{+1.1}_{-0.9} \pm 0.5$		1,2 ABE	02k	BELL	Repl. by WANG 04		
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							
<sup>2</sup> Explicitly vetoes resonant production of $p\overline{p}$ from Charmonium states.							
<sup>3</sup> Provides also results with $m_{p\overline{p}} < 2.85 \text{ GeV}/c^2$ and angular asymmetry of $p\overline{p}$ system.							

$\Gamma(p\overline{p}K^+)/\Gamma(J/\psi(1S)K^+)$				$\Gamma_{463}/\Gamma_{255}$			
VALUE		DOCUMENT ID	TECN	COMMENT			
<b><math>0.0104 \pm 0.0005 \pm 0.0001</math></b>		1,2 AAIJ	13s	LHCB	$p\overline{p}$ at 7 TeV		
1 AAIJ 13s reports $[\Gamma(B^+ \rightarrow p\overline{p}K^+)/\Gamma(B^+ \rightarrow J/\psi(1S)K^+)] / [B(J/\psi(1S) \rightarrow p\overline{p})] = 4.91 \pm 0.19 \pm 0.14$ which we multiply by our best value $B(J/\psi(1S) \rightarrow p\overline{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.							
2 Measurement includes contribution where $p\overline{p}$ is produced in charmonia decays.							

$\Gamma(\Theta(1710)^{++}\overline{p}, \Theta^{++} \rightarrow pK^+)/\Gamma_{\text{total}}$				$\Gamma_{464}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;0.091</b>	90	1 WANG	05A	BELL	$e^+e^- \rightarrow \gamma(4S)$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
<0.1	90	1,2 AUBERT,B	05L	BABR	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							
2 Provides upper limits depending on the pentaquark masses between 1.43 to 2.0 GeV/ $c^2$ .							

$\Gamma(f_J(2220)K^+, f_J \rightarrow p\overline{p})/\Gamma_{\text{total}}$				$\Gamma_{465}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;0.41</b>	90	1 WANG	05A	BELL	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(p\overline{\Lambda}(1520))/\Gamma_{\text{total}}$				$\Gamma_{466}/\Gamma$			
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b><math>3.15 \pm 0.48 \pm 0.27</math></b>		1 AAIJ	14AF	LHCB	$p\overline{p}$ at 7, 8 TeV		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
$3.9^{+1.0}_{-0.9} \pm 0.3$		1 AAIJ	13AU	LHCB	Repl. by AAIJ 14AF		
<15	90	2 AUBERT,B	05L	BABR	$e^+e^- \rightarrow \gamma(4S)$		
1 Uses $B(B^+ \rightarrow J/\psi K^+) = (1.016 \pm 0.033) \times 10^{-3}$ , $B(J/\psi \rightarrow p\overline{p}) = (2.17 \pm 0.07) \times 10^{-3}$ and $B(\Lambda(1520) \rightarrow K^- p) = 0.234 \pm 0.016$ .							
2 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(p\overline{p}K^* \text{ nonresonant})/\Gamma_{\text{total}}$				$\Gamma_{467}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;89</b>	90	BERGFELD	96B	CLE2	$e^+e^- \rightarrow \gamma(4S)$		

$(p\overline{p}K^*(892)^+)/\Gamma_{\text{total}}$			$\Gamma_{468}/\Gamma$		
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>3.6^{+0.8}_{-0.7}</math></b>	<b>OUR AVERAGE</b>				
$3.38^{+0.73}_{-0.60} \pm 0.39$		1,2 CHEN	08c	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$5.3 \pm 1.5 \pm 1.3$		2 AUBERT	07AV	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$10.3^{+3.6}_{-2.8}^{+1.3}_{-1.7}$		2,3 WANG	04	BELL	Repl. by CHEN 08c
1 Explicitly vetoes resonant production of $p\overline{p}$ from charmonium states.					
2 Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
3 Explicitly vetoes resonant production of $p\overline{p}$ from charmonium states. The branching fraction for $M_{p\overline{p}} < 2.85 \text{ GeV}/c^2$ is also reported.					

$\Gamma(f_J(2220)K^{*+}, f_J \rightarrow p\overline{p})/\Gamma_{\text{total}}$				$\Gamma_{469}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;0.77</b>	90	1 AUBERT	07AV	BABR	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(p\overline{\Lambda})/\Gamma_{\text{total}}$				$\Gamma_{470}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt; 0.32</b>	90	1 TSAI	07	BELL	$e^+e^- \rightarrow \gamma(4S)$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
< 0.49	90	1 CHANG	05	BELL	Repl. by TSAI 07		
< 1.5	90	1 BORNHEIM	03	CLE2	$e^+e^- \rightarrow \gamma(4S)$		
< 2.2	90	1 ABE	02o	BELL	$e^+e^- \rightarrow \gamma(4S)$		
< 2.6	90	1 COAN	99	CLE2	$e^+e^- \rightarrow \gamma(4S)$		
<60	90	2 AVERY	89B	CLEO	$e^+e^- \rightarrow \gamma(4S)$		
<93	90	3 ALBRECHT	88F	ARG	$e^+e^- \rightarrow \gamma(4S)$		

- 1 Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
2 AVERY 89B reports  $< 5 \times 10^{-5}$  assuming the  $\gamma(4S)$  decays 43% to  $B^0\overline{B}^0$ . We rescale to 50%.  
3 ALBRECHT 88F reports  $< 8.5 \times 10^{-5}$  assuming the  $\gamma(4S)$  decays 45% to  $B^0\overline{B}^0$ . We rescale to 50%.

$\Gamma(p\overline{\Lambda}\gamma)/\Gamma_{\text{total}}$				$\Gamma_{471}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b><math>2.45^{+0.44}_{-0.38} \pm 0.22</math></b>		1 WANG	07c	BELL	$e^+e^- \rightarrow \gamma(4S)$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
$2.16^{+0.58}_{-0.53} \pm 0.20$		1 LEE	05	BELL	Repl. by WANG 07c		
<3.9	90	2 EDWARDS	03	CLE2	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							
2 Corresponds to $E_\gamma > 1.5 \text{ GeV}$ . The limit changes to $3.3 \times 10^{-6}$ for $E_\gamma > 2.0 \text{ GeV}$ .							

$\Gamma(p\overline{\Lambda}\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{472}/\Gamma$			
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT			
<b><math>3.00^{+0.61}_{-0.53} \pm 0.33</math></b>		1 WANG	07c	BELL	$e^+e^- \rightarrow \gamma(4S)$		

- 1 Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(p\overline{\Sigma}(1385)^0)/\Gamma_{\text{total}}$				$\Gamma_{473}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;0.47</b>	90	1 WANG	07c	BELL	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(\Delta^+\overline{\Lambda})/\Gamma_{\text{total}}$				$\Gamma_{474}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;0.82</b>	90	1 WANG	07c	BELL	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(p\overline{\Sigma}\gamma)/\Gamma_{\text{total}}$				$\Gamma_{475}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;4.6</b>	90	1 LEE	05	BELL	$e^+e^- \rightarrow \gamma(4S)$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
<7.9	90	2 EDWARDS	03	CLE2	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							
2 Corresponds to $E_\gamma > 1.5 \text{ GeV}$ . The limit changes to $6.4 \times 10^{-6}$ for $E_\gamma > 2.0 \text{ GeV}$ .							

$\Gamma(p\overline{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{476}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b><math>5.92^{+0.88}_{-0.84} \pm 0.69</math></b>		1 CHEN	09c	BELL	$e^+e^- \rightarrow \gamma(4S)$		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
<200	90	2 ALBRECHT	88F	ARG	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							
2 ALBRECHT 88F reports $< 1.8 \times 10^{-4}$ assuming the $\gamma(4S)$ decays 45% to $B^0\overline{B}^0$ . We rescale to 50%.							

$\Gamma(p\overline{\Lambda}\rho^0)/\Gamma_{\text{total}}$				$\Gamma_{477}/\Gamma$			
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT			
<b><math>4.78^{+0.67}_{-0.64} \pm 0.60</math></b>		1 CHEN	09c	BELL	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(p\overline{\Lambda}\eta(1270))/\Gamma_{\text{total}}$				$\Gamma_{478}/\Gamma$			
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT			
<b><math>2.03^{+0.77}_{-0.72} \pm 0.27</math></b>		1 CHEN	09c	BELL	$e^+e^- \rightarrow \gamma(4S)$		
1 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(\Lambda\overline{\Lambda}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{479}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT			
<b>&lt;0.94</b>	90	1,2 CHANG	09	BELL	Repl. by CHANG 09		
• • • We do not use the following data for averages, fits, limits, etc. • • •							
<2.8	90	2 LEE	04	BELL	$e^+e^- \rightarrow \gamma(4S)$		
1 For $m_{\Lambda\overline{\Lambda}} < 2.85 \text{ GeV}/c^2$ .							
2 Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .							

$\Gamma(\Lambda\overline{\Lambda}K^+)/\Gamma_{\text{total}}$				$\Gamma_{480}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>3.38^{+0.41}_{-0.36} \pm 0.41</math></b>	<sup>1,2</sup> CHANG	09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.91^{+0.9}_{-0.70} \pm 0.38$	<sup>2</sup> LEE	04	BELL	Repl. by CHANG 09
<sup>1</sup> Excluding charmonium events in $2.85 < m_{\Lambda\overline{\Lambda}K} < 3.128 \text{ GeV}/c^2$ and $3.315 < m_{\Lambda\overline{\Lambda}K} < 3.735 \text{ GeV}/c^2$ . Measurements in various $m_{\Lambda\overline{\Lambda}K}$ bins are also reported.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\Lambda\bar{\Lambda}K^+)/\Gamma_{\text{total}}$				$\Gamma_{481}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	

**$2.19 \pm 1.13$   
 $-0.88 \pm 0.33$**  <sup>1,2</sup> CHANG 09 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> For  $m_{\Lambda\bar{\Lambda}} < 2.85 \text{ GeV}/c^2$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0\rho)/\Gamma_{\text{total}}$				$\Gamma_{482}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$< 1.38$**  90 <sup>1</sup> WEI 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

**$< 380$**  90 <sup>2</sup> BOROLETTO89 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> BOROLETTO 89 reports  $< 3.3 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(\Delta^{++}\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{483}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$< 0.14$**  90 <sup>1</sup> WEI 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

**$< 150$**  90 <sup>2</sup> BOROLETTO89 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> BOROLETTO 89 reports  $< 1.3 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(D^+p\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{484}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**$< 1.5 \times 10^{-5}$**  90 <sup>1</sup> ABE 02W BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^*(2010)^+p\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{485}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**$< 1.5 \times 10^{-5}$**  90 <sup>1</sup> ABE 02W BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0\rho\bar{p}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{486}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	

**$3.72 \pm 0.11 \pm 0.25$**  <sup>1,2</sup> DEL-AMO-SA..12 BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the values of  $D$  and  $D^*$  branching fractions from PDG 08.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^{*0}\rho\bar{p}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{487}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	

**$3.73 \pm 0.17 \pm 0.27$**  <sup>1,2</sup> DEL-AMO-SA..12 BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the values of  $D$  and  $D^*$  branching fractions from PDG 08.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^-p\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{488}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	

**$1.66 \pm 0.13 \pm 0.27$**  <sup>1,2</sup> DEL-AMO-SA..12 BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the values of  $D$  and  $D^*$  branching fractions from PDG 08.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^{*-}p\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{489}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	

**$1.86 \pm 0.16 \pm 0.19$**  <sup>1,2</sup> DEL-AMO-SA..12 BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the values of  $D$  and  $D^*$  branching fractions from PDG 08.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(p\bar{\Lambda}^0\bar{D}^0)/\Gamma_{\text{total}}$				$\Gamma_{490}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT	

**$1.43 \pm 0.25 \pm 0.18$**  <sup>1,2</sup> CHEN 11F BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = 63.9 \pm 0.5\%$ ,  $B(D^0 \rightarrow K^-\pi^+) = 3.89 \pm 0.05\%$ , and  $B(D^0 \rightarrow K^-\pi^+\pi^0) = 13.9 \pm 0.5\%$ .

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  from Upsilon(4S) decays.

$\Gamma(p\bar{\Lambda}^0\bar{D}^*(2007)^0)/\Gamma_{\text{total}}$				$\Gamma_{491}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$< 5$**  90 <sup>1,2,3</sup> CHEN 11F BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> CHEN 11F reports  $< 4.8 \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow p\bar{\Lambda}^0\bar{D}^*(2007)^0)/\Gamma_{\text{total}}]/[B(D^*(2007)^0 \rightarrow D^0\pi^0)]$  assuming  $B(D^*(2007)^0 \rightarrow D^0\pi^0) = (61.9 \pm 2.9) \times 10^{-2}$ , which we rescale to our best value  $B(D^*(2007)^0 \rightarrow D^0\pi^0) = 64.7 \times 10^{-2}$ .

<sup>2</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = 63.9 \pm 0.5\%$  and  $B(D^0 \rightarrow K^-\pi^+) = 3.89 \pm 0.05\%$ .

<sup>3</sup> Assumes equal production of  $B^0$  and  $B^+$  from Upsilon(4S) decays.

$\Gamma(\bar{\Lambda}_c^-p\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{492}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	

**$2.2 \pm 0.4$  OUR AVERAGE** Error includes scale factor of 2.2. See the ideogram below.

**$2.68 \pm 0.15 \pm 0.14$**  <sup>1,2</sup> AUBERT 08BN BABR  $e^+e^- \rightarrow \Upsilon(4S)$

**$1.58 \pm 0.20 \pm 0.08$**  <sup>1,3</sup> GABYSHEV 06A BELL  $e^+e^- \rightarrow \Upsilon(4S)$

**$1.9 \pm 0.5 \pm 0.1$**  <sup>1,4</sup> DYTMAN 02 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

**$1.5 \pm 0.4 \pm 0.1$**  <sup>1,5</sup> GABYSHEV 02 BELL Repl. by GABYSHEV 06A

**$6.2 \pm 2.3$   
 $-2.0 \pm 1.6$**  <sup>1,6</sup> FU 97 CLE2 Repl. by DYTMAN 02

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> AUBERT 08BN reports  $(3.4 \pm 0.1 \pm 0.9) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{\Lambda}_c^-p\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

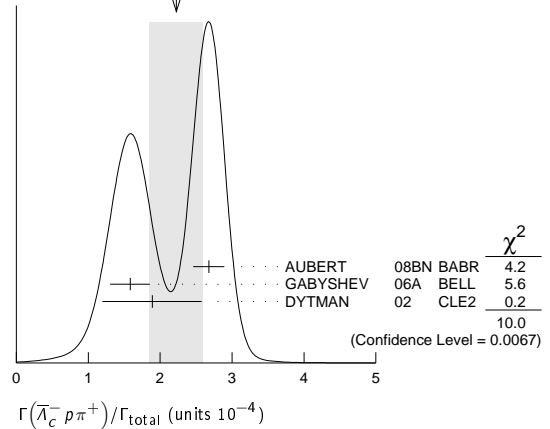
<sup>3</sup> GABYSHEV 06A reports  $(2.01 \pm 0.15 \pm 0.20) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{\Lambda}_c^-p\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> DYTMAN 02 reports  $(2.4 \pm 0.63 \pm 0.62) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{\Lambda}_c^-p\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>5</sup> GABYSHEV 02 reports  $(1.87 \pm 0.51 \pm 0.49) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{\Lambda}_c^-p\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>6</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching fraction.

WEIGHTED AVERAGE  
 $2.2 \pm 0.4$  (Error scaled by 2.2)



$\Gamma(\bar{\Lambda}_c^- \Delta(1232)^{++})/\Gamma_{\text{total}}$				$\Gamma_{493}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$< 1.9$**  90 GABYSHEV 06A BELL  $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\bar{\Lambda}_c^- \Delta_X(1600)^{++})/\Gamma_{\text{total}}$				$\Gamma_{494}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT	

**$4.6 \pm 0.9 \pm 0.2$**  <sup>1</sup> GABYSHEV 06A BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> GABYSHEV 06A reports  $(5.9 \pm 1.0 \pm 0.6) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{\Lambda}_c^- \Delta_X(1600)^{++})/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{\Lambda}_c^- \Delta_X(2420)^{++})/\Gamma_{\text{total}}$				$\Gamma_{495}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT	

**$3.7 \pm 0.8 \pm 0.2$**  <sup>1</sup> GABYSHEV 06A BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> GABYSHEV 06A reports  $(4.7 \pm 1.0 \pm 0.9) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \bar{\Lambda}_c^- \Delta_X(2420)^{++})/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

$B^\pm$

## $\Gamma(\overline{(\Lambda_c^- p)_s \pi^+})/\Gamma_{\text{total}}$ $\Gamma_{496}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.1^{+0.7}_{-0.6} \pm 0.2</math></b>	<sup>1</sup> GABYSHEV	06A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> GABYSHEV 06A reports  $(3.9^{+0.8}_{-0.7} \pm 0.4) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow (\overline{\Lambda_c^- p)_s \pi^+})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.05$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{\Sigma_c(2520)^0 p})/\Gamma_{\text{total}}$ $\Gamma_{497}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.3</b>	90	<sup>1,2</sup> AUBERT	08BN BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<2.7 90 <sup>1,2</sup> GABYSHEV 06A BELL  $e^+ e^- \rightarrow \Upsilon(4S)$   
<4.6 90 <sup>1,2</sup> GABYSHEV 02 BELL Repl. by GABYSHEV 06A  
<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses the value for  $\Lambda_c \rightarrow p K^- \pi^+$  branching ratio  $(5.0 \pm 1.3)\%$ .

## $\Gamma(\overline{\Sigma_c(2520)^0 p})/\Gamma(\overline{\Lambda_c^- p \pi^+})$ $\Gamma_{497}/\Gamma_{492}$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;9</b>	90	AUBERT	08BN BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $\Gamma(\overline{\Sigma_c(2800)^0 p})/\Gamma_{\text{total}}$ $\Gamma_{498}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.6 \pm 0.7 \pm 0.4</math></b>	<sup>1</sup> AUBERT	08BN BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 08BN reports  $[\Gamma(B^+ \rightarrow \overline{\Sigma_c(2800)^0 p})/\Gamma_{\text{total}}] / [\text{B}(B^+ \rightarrow \overline{\Lambda_c^- p \pi^+})] = 0.117 \pm 0.023 \pm 0.024$  which we multiply by our best value  $\text{B}(B^+ \rightarrow \overline{\Lambda_c^- p \pi^+}) = (2.2 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{\Lambda_c^- p \pi^+ \pi^0})/\Gamma_{\text{total}}$ $\Gamma_{499}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.81 \pm 0.29^{+0.52}_{-0.50}</math></b>		<sup>1,2</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<3.12 90 <sup>3</sup> FU 97 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$   
<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> DYTMAN 02 measurement uses  $\text{B}(\Lambda_c^- \rightarrow \overline{p} K^+ \pi^-) = 5.0 \pm 1.3\%$ . The second error includes the systematic and the uncertainty of the branching ratio.  
<sup>3</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

## $\Gamma(\overline{\Lambda_c^- p \pi^+ \pi^+ \pi^- \pi^0})/\Gamma_{\text{total}}$ $\Gamma_{500}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.25 \pm 0.25^{+0.63}_{-0.61}</math></b>		<sup>1,2</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<1.46 90 <sup>3</sup> FU 97 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$   
<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> DYTMAN 02 measurement uses  $\text{B}(\Lambda_c^- \rightarrow \overline{p} K^+ \pi^-) = 5.0 \pm 1.3\%$ . The second error includes the systematic and the uncertainty of the branching ratio.  
<sup>3</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

## $\Gamma(\overline{\Lambda_c^- p \pi^+ \pi^+ \pi^- \pi^0})/\Gamma_{\text{total}}$ $\Gamma_{501}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.34 \times 10^{-2}</math></b>	90	<sup>1</sup> FU	97	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

## $\Gamma(\Lambda_c^+ \Lambda_c^- K^+)/\Gamma_{\text{total}}$ $\Gamma_{502}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>6.9 \pm 2.2</math> OUR AVERAGE</b>			
$9.0 \pm 4.4 \pm 0.5$	<sup>1,2</sup> AUBERT	08H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$6.2^{+2.5}_{-2.4} \pm 0.3$	<sup>2,3</sup> GABYSHEV	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 08H reports  $(1.14 \pm 0.15 \pm 0.62) \times 10^{-3}$  from a measurement of  $[\Gamma(B^+ \rightarrow \Lambda_c^+ \Lambda_c^- K^+)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> GABYSHEV 06 reports  $(7.9^{+1.0}_{-0.9} \pm 3.6) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \Lambda_c^+ \Lambda_c^- K^+)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{\Sigma_c(2455)^0 p})/\Gamma_{\text{total}}$ $\Gamma_{503}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.9 \pm 0.6^{+0.2}_{-0.1}</math></b>		<sup>1,2</sup> GABYSHEV	06A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<8 90 <sup>1,3</sup> DYTMAN 02 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$   
<9.3 90 <sup>1,4</sup> GABYSHEV 02 BELL Repl. by GABYSHEV 06A  
<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> GABYSHEV 06A reports  $(3.7 \pm 0.7 \pm 0.4) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \overline{\Sigma_c(2455)^0 p})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.05$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> DYTMAN 02 measurement uses  $\text{B}(\Lambda_c^- \rightarrow \overline{p} K^+ \pi^-) = 5.0 \pm 1.3\%$ . The second error includes the systematic and the uncertainty of the branching ratio.  
<sup>4</sup> Uses the value for  $\Lambda_c \rightarrow p K^- \pi^+$  branching ratio  $(5.0 \pm 1.3)\%$ .

## $\Gamma(\overline{\Sigma_c(2455)^0 p})/\Gamma(\overline{\Lambda_c^- p \pi^+})$ $\Gamma_{503}/\Gamma_{492}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.123 \pm 0.012 \pm 0.008</math></b>	<sup>1</sup> AUBERT	08BN BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $\Gamma(\overline{\Sigma_c(2455)^0 p \pi^0})/\Gamma_{\text{total}}$ $\Gamma_{504}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.5 \pm 1.1 \pm 0.2</math></b>	<sup>1,2</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> DYTMAN 02 reports  $(4.4 \pm 1.4) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \overline{\Sigma_c(2455)^0 p \pi^0})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.05$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\overline{\Sigma_c(2455)^0 p \pi^- \pi^+})/\Gamma_{\text{total}}$ $\Gamma_{505}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.5 \pm 1.0 \pm 0.2</math></b>	<sup>1,2</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> DYTMAN 02 reports  $(4.4 \pm 1.3) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \overline{\Sigma_c(2455)^0 p \pi^- \pi^+})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.05$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\overline{\Sigma_c(2455)^- p \pi^+ \pi^+})/\Gamma_{\text{total}}$ $\Gamma_{506}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.34 \pm 0.20</math> OUR AVERAGE</b>			
$2.35 \pm 0.16^{+0.13}_{-0.12}$	<sup>1,2</sup> LEES	12Z	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$2.2 \pm 0.8 \pm 0.1$	<sup>1,3</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> LEES 12Z reports  $(2.98 \pm 0.16 \pm 0.15 \pm 0.77) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \overline{\Sigma_c(2455)^- p \pi^+ \pi^+})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> DYTMAN 02 reports  $(2.8 \pm 0.9 \pm 0.5 \pm 0.7) \times 10^{-4}$  from a measurement of  $[\Gamma(B^+ \rightarrow \overline{\Sigma_c(2455)^- p \pi^+ \pi^+})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{\Lambda_c(2593)^-}/\Lambda_c(2625)^- p \pi^+)/\Gamma_{\text{total}}$ $\Gamma_{507}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.9 \times 10^{-4}</math></b>	90	<sup>1,2</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> DYTMAN 02 measurement uses  $\text{B}(\Lambda_c^- \rightarrow \overline{p} K^+ \pi^-) = 5.0 \pm 1.3\%$ . The second error includes the systematic and the uncertainty of the branching ratio.

## $\Gamma(\Xi^0 \Lambda_c^+, \Xi^0 \rightarrow \Xi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{508}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.4 \pm 0.9</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.		
$2.0 \pm 0.7 \pm 0.1$	<sup>1,2</sup> AUBERT	08H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$4.4^{+1.8}_{-1.5} \pm 0.2$	<sup>2,3</sup> CHISTOV	06A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 08H reports  $(2.51 \pm 0.89 \pm 0.61) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \Xi^0 \Lambda_c^+, \Xi^0 \rightarrow \Xi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> CHISTOV 06A reports  $(5.6^{+1.9}_{-1.5} \pm 1.9) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \Xi_c^0 \Lambda_c^+, \Xi_c^0 \rightarrow \Xi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(\Xi_c^0 \Lambda_c^+, \Xi_c^0 \rightarrow \Lambda K^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{509}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.1 \pm 0.9</math> OUR AVERAGE</b>	Error includes scale factor of 1.5.		
$1.3 \pm 0.8 \pm 0.1$	<sup>1,2</sup> AUBERT	08H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$3.1^{+1.1}_{-0.9} \pm 0.2$	<sup>2,3</sup> CHISTOV	06A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 08H reports  $(1.70 \pm 0.93 \pm 0.53) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \Xi_c^0 \Lambda_c^+, \Xi_c^0 \rightarrow \Lambda K^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> CHISTOV 06A reports  $(4.0^{+1.1}_{-0.9} \pm 1.3) \times 10^{-5}$  from a measurement of  $[\Gamma(B^+ \rightarrow \Xi_c^0 \Lambda_c^+, \Xi_c^0 \rightarrow \Lambda K^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(\pi^+ \ell^+ \ell^-)/\Gamma_{\text{total}}$ $\Gamma_{510}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.9 \times 10^{-8}</math></b>	90	<sup>1</sup> WEI	08A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 6.6 \times 10^{-8}$	90	<sup>1</sup> LEES	13M BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.2 \times 10^{-7}$	90	<sup>1</sup> AUBERT	07AG BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(\pi^+ e^+ e^-)/\Gamma_{\text{total}}$ $\Gamma_{511}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 8.0 \times 10^{-8}</math></b>	90	<sup>1</sup> WEI	08A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 12.5 \times 10^{-8}$	90	<sup>1</sup> LEES	13M BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 18 \times 10^{-8}$	90	<sup>1</sup> AUBERT	07AG BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 3.9 \times 10^{-3}$	90	<sup>2</sup> WEIR	90B MRK2	$e^+ e^-$ 29 GeV

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> WEIR 90B assumes  $B^+$  production cross section from LUND.

### $\Gamma(\pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{512}/\Gamma$

VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.79 \pm 0.22 \pm 0.05</math></b>		<sup>1</sup> AAIJ	15AR LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 5.5$	90	<sup>2</sup> LEES	13M BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$2.3 \pm 0.6 \pm 0.1$		<sup>2</sup> AAIJ	12AY LHCB	Repl. by AAIJ 15AR
$< 6.9$	90	<sup>2</sup> WEI	08A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 28$	90	<sup>2</sup> AUBERT	07AG BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 15AR reports  $(1.83 \pm 0.24 \pm 0.05) \times 10^{-8}$  from a measurement of  $[\Gamma(B^+ \rightarrow \pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}] / [\text{B}(B^+ \rightarrow J/\psi(1S) K^+)] / [\text{B}(J/\psi(1S) \rightarrow \mu^+ \mu^-)]$  assuming  $\text{B}(B^+ \rightarrow J/\psi(1S) K^+) = (1.05 \pm 0.05) \times 10^{-3}$ ,  $\text{B}(J/\psi(1S) \rightarrow \mu^+ \mu^-) = (5.961 \pm 0.033) \times 10^{-2}$ , which we rescale to our best values  $\text{B}(B^+ \rightarrow J/\psi(1S) K^+) = (1.026 \pm 0.031) \times 10^{-3}$ ,  $\text{B}(J/\psi(1S) \rightarrow \mu^+ \mu^-) = (5.961 \pm 0.033) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(\pi^+ \mu^+ \mu^-)/\Gamma(K^+ \mu^+ \mu^-)$ $\Gamma_{512}/\Gamma_{516}$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.053 \pm 0.014 \pm 0.001$	AAIJ	12AY LHCB	Repl. by AAIJ 15AR

### $\Gamma(\pi^+ \nu \bar{\nu})/\Gamma_{\text{total}}$ $\Gamma_{513}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 9.8 \times 10^{-5}</math></b>		<sup>1</sup> LUTZ	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.7 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.0 \times 10^{-4}$	90	<sup>1</sup> AUBERT	05H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(K^+ \ell^+ \ell^-)/\Gamma_{\text{total}}$ $\Gamma_{514}/\Gamma$

Test for  $\Delta B=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.51 \pm 0.23</math> OUR AVERAGE</b>		Error includes scale factor of 1.1.		
$4.36 \pm 0.15 \pm 0.18$		<sup>1</sup> AAIJ	13H LHCB	$pp$ at 7 TeV
$4.8 \pm 0.9 \pm 0.2$		<sup>2</sup> AUBERT	09T BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$5.3^{+0.6}_{-0.5} \pm 0.3$		<sup>2</sup> WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.8^{+0.9}_{-0.8} \pm 0.2$		<sup>2</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$5.3^{+1.1}_{-1.0} \pm 0.3$		<sup>2</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A

<sup>1</sup> Uses  $\text{B}(B^+ \rightarrow J/\psi K^+ \rightarrow \mu^+ \mu^- K^+) = (6.01 \pm 0.21) \times 10^{-5}$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(K^+ e^+ e^-)/\Gamma_{\text{total}}$ $\Gamma_{515}/\Gamma$

Test for  $\Delta B=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.5 \pm 0.7</math> OUR AVERAGE</b>		Error includes scale factor of 1.2.		
$5.1^{+1.2}_{-1.1} \pm 0.2$		<sup>1</sup> AUBERT	09T BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$5.7^{+0.9}_{-0.8} \pm 0.3$		<sup>1</sup> WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.2^{+1.2}_{-1.1} \pm 0.2$		<sup>1</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$10.5^{+2.5}_{-2.2} \pm 0.7$		<sup>1</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$6.3^{+1.9}_{-1.7} \pm 0.3$		<sup>2</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
$< 14$	90	<sup>1</sup> ABE	02L BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 9$	90	<sup>1</sup> AUBERT	02L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 24$	90	<sup>3</sup> ANDERSON	01B CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 990$	90	<sup>4</sup> ALBRECHT	91E ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 68000$	90	<sup>5</sup> WEIR	90B MRK2	$e^+ e^-$ 29 GeV
$< 600$	90	<sup>6</sup> AVERY	89B CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 2500$	90	<sup>7</sup> AVERY	87 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ . The second error is a total of systematic uncertainties including model dependence.

<sup>3</sup> The result is for di-lepton masses above 0.5 GeV.

<sup>4</sup> ALBRECHT 91E reports  $< 9.0 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>5</sup> WEIR 90B assumes  $B^+$  production cross section from LUND.

<sup>6</sup> AVERY 89B reports  $< 5 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>7</sup> AVERY 87 reports  $< 2.1 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

### $\Gamma(K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$ $\Gamma_{516}/\Gamma$

Test for  $\Delta B=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.43 \pm 0.24</math> OUR FIT</b>		Error includes scale factor of 1.2.		
<b><math>4.36 \pm 0.27</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.		
$4.29 \pm 0.07 \pm 0.21$		<sup>1</sup> AAIJ	14M LHCB	$pp$ at 7, 8 TeV
$4.1^{+1.6}_{-1.5} \pm 0.2$		<sup>2</sup> AUBERT	09T BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$5.3^{+0.8}_{-0.7} \pm 0.3$		<sup>2</sup> WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.36 \pm 0.15 \pm 0.18$		<sup>3</sup> AAIJ	13H LHCB	Repl. by AAIJ 14M
$3.1^{+1.5}_{-1.2} \pm 0.3$		<sup>2</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$0.7^{+1.9}_{-1.1} \pm 0.2$		<sup>2</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$4.5^{+1.4}_{-1.2} \pm 0.3$		<sup>4</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
$9.8^{+4.6}_{-3.6} \pm 1.6$		<sup>2</sup> ABE	02L BELL	Repl. by ISHIKAWA 03
$< 12$	90	<sup>2</sup> AUBERT	02L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 36.8$	90	<sup>5</sup> ANDERSON	01B CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 52$	90	<sup>6</sup> AFFOLDER	99B CDF	$p\bar{p}$ at 1.8 TeV
$< 100$	90	<sup>7</sup> ABE	96L CDF	Repl. by AFFOLDER 99B
$< 2400$	90	<sup>8</sup> ALBRECHT	91E ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 64000$	90	<sup>9</sup> WEIR	90B MRK2	$e^+ e^-$ 29 GeV
$< 1700$	90	<sup>10</sup> AVERY	89B CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 3800$	90	<sup>11</sup> AVERY	87 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $\text{B}(B^+ \rightarrow J/\psi(1S) K^+) = (0.998 \pm 0.014 \pm 0.040) \times 10^{-3}$  for normalization.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> Uses  $\text{B}(B^+ \rightarrow J/\psi K^+ \rightarrow \mu^+ \mu^- K^+) = (6.01 \pm 0.21) \times 10^{-5}$ .

<sup>4</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ . The second error is a total of systematic uncertainties including model dependence.

<sup>5</sup> The result is for di-lepton masses above 0.5 GeV.

<sup>6</sup> AFFOLDER 99B measured relative to  $B^+ \rightarrow J/\psi(1S) K^+$ .

<sup>7</sup> ABE 96L measured relative to  $B^+ \rightarrow J/\psi(1S) K^+$  using PDG 94 branching ratios.

<sup>8</sup> ALBRECHT 91E reports  $< 2.2 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>9</sup> WEIR 90B assumes  $B^+$  production cross section from LUND.

<sup>10</sup> AVERY 89B reports  $< 1.5 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>11</sup> AVERY 87 reports  $< 3.2 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

# Meson Particle Listings

$B^\pm$

## $\Gamma(K^+\mu^+\mu^-)/\Gamma(J/\psi(1S)K^+)$

$\Gamma_{516}/\Gamma_{255}$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.431 \pm 0.025</math> OUR FIT</b>				Error includes scale factor of 1.2.
<b><math>0.46 \pm 0.04 \pm 0.02</math></b>		AALTONEN	11A1	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.38 \pm 0.05 \pm 0.02$		AALTONEN	11L	CDF Repl. by AALTONEN 11A1
$0.59 \pm 0.15 \pm 0.03$		AALTONEN	09B	CDF Repl. by AALTONEN 11L

## $\Gamma(K^+\nu\bar{\nu})/\Gamma_{\text{total}}$

$\Gamma_{517}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.6 \times 10^{-5}</math></b>	90	<sup>1,2</sup> LEES	13i	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5.5 \times 10^{-5}$	90	<sup>1</sup> LUTZ	13	BELL $e^+e^- \rightarrow \gamma(4S)$
$<1.3 \times 10^{-5}$	90	<sup>1</sup> DEL-AMO-SAL0q	BABR	Repl. by LEES 13i
$<1.4 \times 10^{-5}$	90	<sup>1</sup> CHEN	07D	BELL $e^+e^- \rightarrow \gamma(4S)$
$<5.2 \times 10^{-5}$	90	<sup>1</sup> AUBERT	05H	BABR $e^+e^- \rightarrow \gamma(4S)$
$<2.4 \times 10^{-4}$	90	<sup>1</sup> BROWDER	01	CLE2 $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ . <sup>2</sup> Also reported a limit $< 3.7 \times 10^{-5}$ at 90% CL obtained using a fully reconstructed hadronic $B$ -tag evnets.				

## $\Gamma(\rho^+\nu\bar{\nu})/\Gamma_{\text{total}}$

$\Gamma_{518}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.13 \times 10^{-4}</math></b>	90	<sup>1</sup> LUTZ	13	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.5 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D	BELL Repl. by LUTZ 13
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

## $\Gamma(K^*(892)^+\ell^+\ell^-)/\Gamma_{\text{total}}$

$\Gamma_{519}/\Gamma$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>10.1 \pm 1.1</math> OUR AVERAGE</b>				Error includes scale factor of 1.1.
$9.24 \pm 0.93 \pm 0.67$		AAIJ	14M	LHCB $pp$ at 7, 8 TeV
$14.0 \pm 4.0 \pm 3.7 \pm 0.9$		<sup>1</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$
$12.4 \pm 2.3 \pm 2.1 \pm 1.3$		<sup>1</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$11.6 \pm 1.9$		<sup>2</sup> AAIJ	12AH	LHCB Repl. by AAIJ 14M
$7.3 \pm 5.0 \pm 4.2 \pm 2.1$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T
$<22$	90	<sup>1</sup> ISHIKAWA	03	BELL $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ . <sup>2</sup> Measured in $B^+ \rightarrow K^*(892)^+\mu^+\mu^-$ decays.				

## $\Gamma(K^*(892)^+e^+e^-)/\Gamma_{\text{total}}$

$\Gamma_{520}/\Gamma$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>15.5 \pm 4.0 \pm 3.1</math> OUR AVERAGE</b>				
$13.8 \pm 4.7 \pm 4.2 \pm 0.8$		<sup>1</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$
$17.3 \pm 5.0 \pm 4.2 \pm 2.0$		<sup>1</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$7.5 \pm 7.6 \pm 6.5 \pm 3.8$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T
$2.0 \pm 13.4 \pm 8.7 \pm 2.8$		<sup>1</sup> AUBERT	03U	BABR $e^+e^- \rightarrow \gamma(4S)$
$< 46$	90	<sup>2</sup> ISHIKAWA	03	BELL $e^+e^- \rightarrow \gamma(4S)$
$< 89$	90	<sup>1</sup> ABE	02	BELL Repl. by ISHIKAWA 03
$< 95$	90	<sup>1</sup> AUBERT	02L	BABR $e^+e^- \rightarrow \gamma(4S)$
$<6900$	90	<sup>3</sup> ALBRECHT	91E	ARG $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ . <sup>2</sup> Assumes equal production of $B^0$ and $B^+$ at $\gamma(4S)$ . The second error is a total of systematic uncertainties including model dependence. <sup>3</sup> ALBRECHT 91E reports $< 6.3 \times 10^{-4}$ assuming the $\gamma(4S)$ decays 45% to $B^0\bar{B}^0$ . We rescale to 50%.				

## $\Gamma(K^*(892)^+\mu^+\mu^-)/\Gamma_{\text{total}}$

$\Gamma_{521}/\Gamma$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>9.6 \pm 1.0</math> OUR FIT</b>				
<b><math>9.6 \pm 1.1</math> OUR AVERAGE</b>				
$9.24 \pm 0.93 \pm 0.67$		<sup>1</sup> AAIJ	14M	LHCB $pp$ at 7, 8 TeV
$14.6 \pm 7.9 \pm 7.5 \pm 1.2$		<sup>2</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$
$11.1 \pm 3.2 \pm 2.7 \pm 1.0$		<sup>2</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$11.6 \pm 1.9$		AAIJ	12AH	LHCB Repl. by AAIJ 14M
$9.7 \pm 9.4 \pm 6.9 \pm 1.4$		<sup>2</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T
$30.7 \pm 25.8 \pm 17.8 \pm 4.2$		<sup>2</sup> AUBERT	03U	BABR $e^+e^- \rightarrow \gamma(4S)$
$6.5 \pm 6.9 \pm 5.3 \pm 1.5 \pm 1.6$		<sup>3</sup> ISHIKAWA	03	BELL Repl. by WEI 09A
$< 39$	90	<sup>2</sup> ABE	02	BELL Repl. by ISHIKAWA 03
$<170$	90	<sup>2</sup> AUBERT	02L	BABR $e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Uses  $B(B^+ \rightarrow J/\psi(1S)K^*(892)^+) = (1.431 \pm 0.027 \pm 0.090) \times 10^{-3}$  for normalization.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>3</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\gamma(4S)$ . The second error is a total of systematic uncertainties including model dependence. The 90% C.L. upper limit is  $2.2 \times 10^{-6}$ .

## $\Gamma(K^*(892)^+\mu^+\mu^-)/\Gamma(J/\psi(1S)K^*(892)^+)$

$\Gamma_{521}/\Gamma_{260}$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.67 \pm 0.08</math> OUR FIT</b>				
<b><math>0.67 \pm 0.22 \pm 0.04</math></b>		AALTONEN	11A1	CDF $p\bar{p}$ at 1.96 TeV

## $\Gamma(K^*(892)^+\nu\bar{\nu})/\Gamma_{\text{total}}$

$\Gamma_{522}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;4.0 \times 10^{-5}</math></b>	90	<sup>1</sup> LUTZ	13	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<6.4 \times 10^{-5}$	90	<sup>1,2</sup> LEES	13i	BABR $e^+e^- \rightarrow \gamma(4S)$
$<8 \times 10^{-5}$	90	AUBERT	08Bc	BABR Repl. by LEES 13i
$<1.4 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D	BELL $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ . <sup>2</sup> Also reported a limit $< 11.6 \times 10^{-5}$ at 90% CL obtained using a fully reconstructed hadronic $B$ -tag evnets.				

## $\Gamma(K^+\pi^+\pi^-\mu^+\mu^-)/\Gamma(\psi(2S)K^+)$

$\Gamma_{523}/\Gamma_{283}$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>6.95 \pm 0.46 \pm 0.43 \pm 0.34</math></b>		AAIJ	14Az	LHCB $pp$ at 7, 8 TeV

## $\Gamma(\phi K^+\mu^+\mu^-)/\Gamma(J/\psi(1S)\phi K^+)$

$\Gamma_{524}/\Gamma_{267}$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.58 \pm 0.36 \pm 0.19 \pm 0.32 \pm 0.07</math></b>		AAIJ	14Az	LHCB $pp$ at 7, 8 TeV

## $\Gamma(\pi^+e^+\mu^-)/\Gamma_{\text{total}}$

$\Gamma_{525}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.0064</math></b>	90	<sup>1</sup> WEIR	90B	MRK2 $e^+e^-$ 29 GeV
<sup>1</sup> WEIR 90B assumes $B^+$ production cross section from LUND.				

## $\Gamma(\pi^+e^-\mu^+)/\Gamma_{\text{total}}$

$\Gamma_{526}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.0064</math></b>	90	<sup>1</sup> WEIR	90B	MRK2 $e^+e^-$ 29 GeV
<sup>1</sup> WEIR 90B assumes $B^+$ production cross section from LUND.				

## $\Gamma(\pi^+e^\pm\tau^\mp)/\Gamma_{\text{total}}$

$\Gamma_{527}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.7 \times 10^{-7}</math></b>	90	<sup>1</sup> AUBERT	07Ag	BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

## $\Gamma(\pi^+\tau^+\tau^-)/\Gamma_{\text{total}}$

$\Gamma_{528}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;74</math></b>	90	<sup>1</sup> LEES	12P	BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.				

## $\Gamma(\pi^+e^-\tau^+)/\Gamma_{\text{total}}$

$\Gamma_{529}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;20</math></b>	90	<sup>1</sup> LEES	12P	BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.				

## $\Gamma(\pi^+e^\pm\tau^\mp)/\Gamma_{\text{total}}$

$\Gamma_{530}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;75</math></b>	90	<sup>1,2</sup> LEES	12P	BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes $B(B^+ \rightarrow h^+\ell^+\tau^-) = B(B^+ \rightarrow h^+\ell^-\tau^+)$ . <sup>2</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.				

## $\Gamma(\pi^+\mu^+\tau^-)/\Gamma_{\text{total}}$

$\Gamma_{531}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;62</math></b>	90	<sup>1</sup> LEES	12P	BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.				

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Meson Particle Listings

$B^\pm$

$\Gamma(\pi^+\mu^-\tau^+)/\Gamma_{\text{total}}$					$\Gamma_{532}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<45	90	<sup>1</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(\pi^+\mu^\pm\tau^\mp)/\Gamma_{\text{total}}$					$\Gamma_{533}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<72	90	<sup>1,2</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes $B(B^+ \rightarrow h^+\ell^+\tau^-) = B(B^+ \rightarrow h^+\ell^-\tau^+)$ .					
<sup>2</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(K^+e^+\mu^-)/\Gamma_{\text{total}}$					$\Gamma_{534}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.91	90	<sup>1</sup> AUBERT,B	06j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
<8	90	<sup>1</sup> AUBERT	02L BABR	Repl. by AUBERT,B 06j	
$<6.4 \times 10^4$	90	<sup>2</sup> WEIR	90B MRK2	$e^+e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(K^+e^-\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{535}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.3	90	<sup>1</sup> AUBERT,B	06j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
$<6.4 \times 10^4$	90	<sup>2</sup> WEIR	90B MRK2	$e^+e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(K^+e^\pm\mu^\mp)/\Gamma_{\text{total}}$					$\Gamma_{536}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.91	90	<sup>1</sup> AUBERT,B	06j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+e^+\tau^-)/\Gamma_{\text{total}}$					$\Gamma_{537}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<43	90	<sup>1</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(K^+e^-\tau^+)/\Gamma_{\text{total}}$					$\Gamma_{538}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<15	90	<sup>1</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(K^+e^\pm\tau^\mp)/\Gamma_{\text{total}}$					$\Gamma_{539}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<30	90	<sup>1,2</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes $B(B^+ \rightarrow h^+\ell^+\tau^-) = B(B^+ \rightarrow h^+\ell^-\tau^+)$ .					
<sup>2</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(K^+\mu^+\tau^-)/\Gamma_{\text{total}}$					$\Gamma_{540}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<45	90	<sup>1</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(K^+\mu^-\tau^+)/\Gamma_{\text{total}}$					$\Gamma_{541}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<28	90	<sup>1</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					

$\Gamma(K^+\mu^\pm\tau^\mp)/\Gamma_{\text{total}}$					$\Gamma_{542}/\Gamma$
Test of lepton family number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<48	90	<sup>1,2</sup> LEES	12P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
<77	90	<sup>1</sup> AUBERT	07AZ BABR	Repl. by LEES 12P	
<sup>1</sup> Uses a fully reconstructed hadronic $B$ decay as a tag on the recoil side.					
<sup>2</sup> Assumes $B(B^+ \rightarrow h^+\ell^+\tau^-) = B(B^+ \rightarrow h^+\ell^-\tau^+)$ .					

$\Gamma(K^*(892)^+e^+\mu^-)/\Gamma_{\text{total}}$					$\Gamma_{543}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<13	90	<sup>1</sup> AUBERT,B	06j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(892)^+e^-\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{544}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<9.9	90	<sup>1</sup> AUBERT,B	06j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(892)^+e^\pm\mu^\mp)/\Gamma_{\text{total}}$					$\Gamma_{545}/\Gamma$
Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<1.4 $\times 10^{-6}$	90	<sup>1</sup> AUBERT,B	06j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
$<7.9 \times 10^{-6}$	90	<sup>1</sup> AUBERT	02L BABR	Repl. by AUBERT,B 06j	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\pi^-e^+e^+)/\Gamma_{\text{total}}$					$\Gamma_{546}/\Gamma$
Test of total lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<2.3 $\times 10^{-8}$	90	<sup>1</sup> LEES	12j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
$<1.6 \times 10^{-6}$	90	<sup>1</sup> EDWARDS	02B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<0.0039	90	<sup>2</sup> WEIR	90B MRK2	$e^+e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(\pi^-\mu^+\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{547}/\Gamma$
Test of total lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
< 4.0 $\times 10^{-9}$	95	<sup>1</sup> AAIJ	14Ac LHCb	$pp$ at 7, 8 TeV	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
$< 1.3 \times 10^{-8}$	95	<sup>2</sup> AAIJ	12Ad LHCb	Repl. by AAIJ 14Ac	
$< 4.4 \times 10^{-8}$	90	<sup>3</sup> AAIJ	12c LHCb	$pp$ at 7 TeV	
$<10.7 \times 10^{-8}$	90	<sup>3</sup> LEES	12j BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
$< 1.4 \times 10^{-6}$	90	<sup>3</sup> EDWARDS	02B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
$< 9.1 \times 10^{-3}$	90	<sup>4</sup> WEIR	90B MRK2	$e^+e^-$ 29 GeV	
<sup>1</sup> Uses $B^+ \rightarrow J/\psi K^+$ , $J/\psi \rightarrow \mu^+\mu^-$ mode for normalization. Obtains neutrino-mass-dependent upper limits in the range $0.4\text{--}4.0 \times 10^{-9}$ . This limit is applicable for Majorana neutrino lifetime $< 1$ ps.					
<sup>2</sup> Uses $B^+ \rightarrow J/\psi K^+$ , $J/\psi \rightarrow \mu^+\mu^-$ mode for normalization. Obtains neutrino-mass-dependent upper limits in the range $0.4\text{--}1.0 \times 10^{-8}$ .					
<sup>3</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>4</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(\pi^-e^+\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{548}/\Gamma$
Test of total lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<1.5 $\times 10^{-7}$	90	<sup>1</sup> LEES	14A BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
$<1.3 \times 10^{-6}$	90	<sup>1</sup> EDWARDS	02B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<0.0064	90	<sup>2</sup> WEIR	90B MRK2	$e^+e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(\rho^-e^+e^+)/\Gamma_{\text{total}}$					$\Gamma_{549}/\Gamma$
Test of total lepton number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.17	90	<sup>1</sup> LEES	14A BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
<2.6	90	<sup>1</sup> EDWARDS	02B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\rho^-\mu^+\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{550}/\Gamma$
Test of total lepton number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.42	90	<sup>1</sup> LEES	14A BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
<5.0	90	<sup>1</sup> EDWARDS	02B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\rho^-e^+\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{551}/\Gamma$
Test of total lepton number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.47	90	<sup>1</sup> LEES	14A BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • •		We do not use the following data for averages, fits, limits, etc. • • •			
<3.3	90	<sup>1</sup> EDWARDS	02B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

# Meson Particle Listings

$B^\pm$

$\Gamma(K^- e^+ e^+)/\Gamma_{\text{total}}$					$\Gamma_{552}/\Gamma$
Test of total lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.0 <math>\times 10^{-8}</math></b>	90	<sup>1</sup> LEES	12j	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<1.0 $\times 10^{-6}$	90	<sup>1</sup> EDWARDS	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
<0.0039	90	<sup>2</sup> WEIR	90B	MRK2 $e^+ e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(K^- \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{553}/\Gamma$
Test of total lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;4.1 <math>\times 10^{-8}</math></b>	90	AALJ	12c	LHCB $pp$ at 7 TeV	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<6.7 $\times 10^{-8}$	90	<sup>1</sup> LEES	12j	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
<1.8 $\times 10^{-6}$	90	<sup>1</sup> EDWARDS	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
<1 $\times 10^{-3}$	90	<sup>2</sup> WEIR	90B	MRK2 $e^+ e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(K^- e^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{554}/\Gamma$
Test of total lepton number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.6 <math>\times 10^{-7}</math></b>	90	<sup>1</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<2.0 $\times 10^{-6}$	90	<sup>1</sup> EDWARDS	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
<0.0064	90	<sup>2</sup> WEIR	90B	MRK2 $e^+ e^-$ 29 GeV	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> WEIR 90B assumes $B^+$ production cross section from LUND.					

$\Gamma(K^*(892)^- e^+ e^+)/\Gamma_{\text{total}}$					$\Gamma_{555}/\Gamma$
Test of total lepton number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.40</b>	90	<sup>1</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<2.8	90	<sup>1</sup> EDWARDS	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(892)^- \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{556}/\Gamma$
Test of total lepton number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.59</b>	90	<sup>1</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<8.3	90	<sup>1</sup> EDWARDS	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(892)^- e^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{557}/\Gamma$
Test of total lepton number conservation.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.30</b>	90	<sup>1</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<4.4	90	<sup>1</sup> EDWARDS	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(D^- e^+ e^+)/\Gamma_{\text{total}}$					$\Gamma_{558}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.6 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
<b>&lt;2.6 <math>\times 10^{-6}</math></b>	90	<sup>1,2</sup> SEON	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^0$ and $B^+$ from Upsilon(4S) decays.					
<sup>2</sup> Uses $D^- \rightarrow K^+ \pi^- \pi^-$ mode and 3-body phase-space hypothesis for the signal decays.					

$\Gamma(D^- e^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{559}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.8 <math>\times 10^{-6}</math></b>	90	<sup>1,2</sup> SEON	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<2.1 $\times 10^{-6}$	90	<sup>1</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^0$ and $B^+$ from Upsilon(4S) decays.					
<sup>2</sup> Uses $D^- \rightarrow K^+ \pi^- \pi^-$ mode and 3-body phase-space hypothesis for the signal decays.					

$\Gamma(D^- \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{560}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 6.9 <math>\times 10^{-7}</math></b>	95	<sup>1</sup> AALJ	12AD	LHCB $pp$ at 7 TeV	
• • • We do not use the following data for averages, fits, limits, etc.		• • •			
<17 $\times 10^{-7}$	90	<sup>2</sup> LEES	14A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
< 1.1 $\times 10^{-6}$	90	<sup>2,3</sup> SEON	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Uses $B^+ \rightarrow \psi(2S) K^+$ , $\psi(2S) \rightarrow J/\psi \pi^+ \pi^-$ mode for normalization.					
<sup>2</sup> Assumes equal production of $B^0$ and $B^+$ from Upsilon(4S) decays.					
<sup>3</sup> Uses $D^- \rightarrow K^+ \pi^- \pi^-$ mode and 3-body phase-space hypothesis for the signal decays.					

$\Gamma(D^{*-} \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{561}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.4 <math>\times 10^{-6}</math></b>	95	<sup>1</sup> AALJ	12AD	LHCB $pp$ at 7 TeV	
<sup>1</sup> Uses $B^+ \rightarrow \psi(2S) K^+$ , $\psi(2S) \rightarrow J/\psi \pi^+ \pi^-$ mode for normalization.					

$\Gamma(D_s^- \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{562}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;5.8 <math>\times 10^{-7}</math></b>	95	<sup>1</sup> AALJ	12AD	LHCB $pp$ at 7 TeV	
<sup>1</sup> Uses $B^+ \rightarrow \psi(2S) K^+$ , $\psi(2S) \rightarrow J/\psi \pi^+ \pi^-$ mode for normalization. Obtains neutrino-mass-dependent upper limits in the range $1.5\text{--}8.0 \times 10^{-7}$ .					

$\Gamma(\overline{D}^0 \pi^- \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{563}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.5 <math>\times 10^{-6}</math></b>	95	<sup>1</sup> AALJ	12AD	LHCB $pp$ at 7 TeV	
<sup>1</sup> Uses $B^+ \rightarrow \psi(2S) K^+$ , $\psi(2S) \rightarrow J/\psi \pi^+ \pi^-$ mode for normalization. Obtains neutrino-mass-dependent upper limits in the range $0.3\text{--}1.5 \times 10^{-6}$ .					

$\Gamma(\Lambda^0 \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{564}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;6 <math>\times 10^{-8}</math></b>	90	<sup>1,2</sup> DEL-AMO-SA...11k	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> DEL-AMO-SANCHEZ 11k reports $< 6.1 \times 10^{-8}$ from a measurement of $[\Gamma(B^+ \rightarrow \Lambda^0 \mu^+)/\Gamma_{\text{total}}] \times [B(\Lambda \rightarrow p \pi^-)]$ assuming $B(\Lambda \rightarrow p \pi^-) = (63.9 \pm 0.5) \times 10^{-2}$ .					
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^0 \overline{B}^0) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^+ B^-) = (48.4 \pm 0.6)\%$ .					

$\Gamma(\Lambda^0 e^+)/\Gamma_{\text{total}}$					$\Gamma_{565}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.2 <math>\times 10^{-8}</math></b>	90	<sup>1,2</sup> DEL-AMO-SA...11k	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> DEL-AMO-SANCHEZ 11k reports $< 3.2 \times 10^{-8}$ from a measurement of $[\Gamma(B^+ \rightarrow \Lambda^0 e^+)/\Gamma_{\text{total}}] \times [B(\Lambda \rightarrow p \pi^-)]$ assuming $B(\Lambda \rightarrow p \pi^-) = (63.9 \pm 0.5) \times 10^{-2}$ .					
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^0 \overline{B}^0) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^+ B^-) = (48.4 \pm 0.6)\%$ .					

$\Gamma(\overline{\Lambda}^0 \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{566}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;6 <math>\times 10^{-8}</math></b>	90	<sup>1,2</sup> DEL-AMO-SA...11k	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> DEL-AMO-SANCHEZ 11k reports $< 6.2 \times 10^{-8}$ from a measurement of $[\Gamma(B^+ \rightarrow \overline{\Lambda}^0 \mu^+)/\Gamma_{\text{total}}] \times [B(\Lambda \rightarrow p \pi^-)]$ assuming $B(\Lambda \rightarrow p \pi^-) = (63.9 \pm 0.5) \times 10^{-2}$ .					
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^0 \overline{B}^0) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^+ B^-) = (48.4 \pm 0.6)\%$ .					

$\Gamma(\overline{\Lambda}^0 e^+)/\Gamma_{\text{total}}$					$\Gamma_{567}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;8 <math>\times 10^{-8}</math></b>	90	<sup>1,2</sup> DEL-AMO-SA...11k	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> DEL-AMO-SANCHEZ 11k reports $< 8.1 \times 10^{-8}$ from a measurement of $[\Gamma(B^+ \rightarrow \overline{\Lambda}^0 e^+)/\Gamma_{\text{total}}] \times [B(\Lambda \rightarrow p \pi^-)]$ assuming $B(\Lambda \rightarrow p \pi^-) = (63.9 \pm 0.5) \times 10^{-2}$ .					
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^0 \overline{B}^0) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^+ B^-) = (48.4 \pm 0.6)\%$ .					

## POLARIZATION IN $B^+$ DECAY

In decays involving two vector mesons, one can distinguish among the states in which meson polarizations are both longitudinal (L) or both are transverse and parallel (||) or perpendicular ( $\perp$ ) to each other with the parameters  $\Gamma_L/\Gamma$ ,  $\Gamma_\perp/\Gamma$ , and the relative phases  $\phi_\parallel$  and  $\phi_\perp$ . See the definitions in the note on “Polarization in  $B$  Decays” review in the  $B^0$  Particle Listings.

$\Gamma_L/\Gamma$ in $B^+ \rightarrow \overline{D}^{*0} \rho^+$					$\Gamma_{568}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.892<math>\pm</math>0.018<math>\pm</math>0.016</b>		CSORNA	03	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
$\Gamma_L/\Gamma$ in $B^+ \rightarrow \overline{D}^{*0} K^{*+}$					$\Gamma_{569}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.86<math>\pm</math>0.06<math>\pm</math>0.03</b>		AUBERT	04K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
$\Gamma_L/\Gamma$ in $B^+ \rightarrow J/\psi K^{*+}$					$\Gamma_{570}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.604<math>\pm</math>0.015<math>\pm</math>0.018</b>		ITOH	05	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
$\Gamma_\perp/\Gamma$ in $B^+ \rightarrow J/\psi K^{*+}$					$\Gamma_{571}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.180<math>\pm</math>0.014<math>\pm</math>0.010</b>		ITOH	05	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
$\Gamma_L/\Gamma$ in $B^+ \rightarrow \omega K^{*+}$					$\Gamma_{572}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.41<math>\pm</math>0.18<math>\pm</math>0.05</b>		AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
$\Gamma_L/\Gamma$ in $B^+ \rightarrow \omega K_2^*(1430)^+$					$\Gamma_{573}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.56<math>\pm</math>0.10<math>\pm</math>0.04</b>		AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	



$\Gamma_L/\Gamma$  in  $B^+ \rightarrow K^{*+} \bar{K}^{*0}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.82 \pm 0.15</math> <math>-0.21</math> OUR AVERAGE</b>			
$1.06 \pm 0.30 \pm 0.14$	<sup>1</sup> GOH	15	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.75 \pm 0.16$ $-0.26 \pm 0.03$	<sup>2,3</sup> AUBERT	09f	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Signal significance 2.7 standard deviations.<sup>2</sup> Signal significance 3.7 standard deviations.<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \phi K^*(892)^+$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.50 \pm 0.05</math> OUR AVERAGE</b>			
$0.49 \pm 0.05 \pm 0.03$	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.52 \pm 0.08 \pm 0.03$	CHEN	05A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.46 \pm 0.12 \pm 0.03$	AUBERT	03v	BABR Repl. by AUBERT 07Ba

 $\Gamma_\perp/\Gamma$  in  $B^+ \rightarrow \phi K^{*+}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.20 \pm 0.05</math> OUR AVERAGE</b>			
$0.21 \pm 0.05 \pm 0.02$	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.19 \pm 0.08 \pm 0.02$	CHEN	05A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

 $\phi_\parallel$  in  $B^+ \rightarrow \phi K^{*+}$ 

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>2.34 \pm 0.18</math> OUR AVERAGE</b>			
$2.47 \pm 0.20 \pm 0.07$	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$2.10 \pm 0.28 \pm 0.04$	CHEN	05A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

 $\phi_\perp$  in  $B^+ \rightarrow \phi K^{*+}$ 

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>2.58 \pm 0.17</math> OUR AVERAGE</b>			
$2.69 \pm 0.20 \pm 0.03$	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$2.31 \pm 0.30 \pm 0.07$	CHEN	05A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

 $\delta_0(B^+ \rightarrow \phi K^{*+})$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>3.07 \pm 0.18 \pm 0.06</math></b>			
	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}^0(B^+ \rightarrow \phi K^{*+})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.17 \pm 0.11 \pm 0.02</math></b>			
	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}^\perp(B^+ \rightarrow \phi K^{*+})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.22 \pm 0.24 \pm 0.08</math></b>			
	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Delta\phi_\parallel(B^+ \rightarrow \phi K^{*+})$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>0.07 \pm 0.20 \pm 0.05</math></b>			
	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Delta\phi_\perp(B^+ \rightarrow \phi K^{*+})$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>0.19 \pm 0.20 \pm 0.07</math></b>			
	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Delta\delta_0(B^+ \rightarrow \phi K^{*+})$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>0.20 \pm 0.18 \pm 0.03</math></b>			
	AUBERT	07Ba	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \phi K_1(1270)^+$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.46 \pm 0.12 \pm 0.06</math> <math>-0.13 - 0.07</math></b>			
	AUBERT	08Bi	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \phi K_2^*(1430)^+$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.80 \pm 0.09</math> <math>-0.10 \pm 0.03</math></b>			
	AUBERT	08Bi	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\delta_0(B^+ \rightarrow \phi K_2^*(1430)^+)$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>3.59 \pm 0.19 \pm 0.12</math></b>			
	AUBERT	08Bi	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Delta\delta_0(B^+ \rightarrow \phi K_2^*(1430)^+)$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>-0.05 \pm 0.19 \pm 0.06</math></b>			
	AUBERT	08Bi	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \rho^0 K^*(892)^+$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.78 \pm 0.12 \pm 0.03</math></b>	DEL-AMO-SA..11D	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.96 \pm 0.04$ $-0.15 \pm 0.04$	AUBERT	03v	BABR Repl. by DEL-AMO-SANCHEZ 11D

 $\Gamma_L/\Gamma(B^+ \rightarrow K^*(892)^0 \rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.48 \pm 0.08</math> OUR AVERAGE</b>			
$0.52 \pm 0.10 \pm 0.04$	AUBERT,B	06G	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.43 \pm 0.11 \pm 0.05$ $-0.02$	ZHANG	05D	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \rho^+ \rho^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.950 \pm 0.016</math> OUR AVERAGE</b>			
$0.950 \pm 0.015 \pm 0.006$	AUBERT	09G	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.948 \pm 0.106 \pm 0.021$	ZHANG	03B	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.905 \pm 0.042 \pm 0.023$ $-0.027$	AUBERT,BE	06G	BABR Repl. by AUBERT 09G
$0.97 \pm 0.03$ $-0.07 \pm 0.04$	AUBERT	03v	BABR Repl. by AUBERT,BE 06G

 $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \omega \rho^+$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.90 \pm 0.05 \pm 0.03</math></b>			
	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.82 \pm 0.11 \pm 0.02$	AUBERT,B	06T	BABR Repl. by AUBERT 09H
$0.88 \pm 0.12 \pm 0.03$ $-0.15$	AUBERT	05o	BABR Repl. by AUBERT,B 06T

 $\Gamma_L/\Gamma$  in  $B^+ \rightarrow \rho^0 K^*(892)^+$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.32 \pm 0.17 \pm 0.09</math></b>			
	CHEN	08c	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

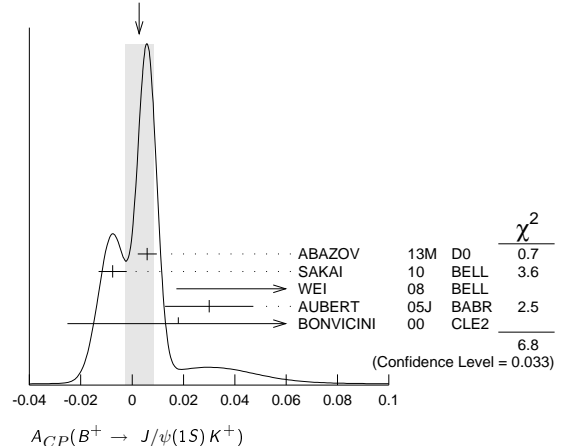
## CP VIOLATION

 $A_{CP}$  is defined as

$$\frac{B(B^- \rightarrow \bar{f}) - B(B^+ \rightarrow f)}{B(B^- \rightarrow \bar{f}) + B(B^+ \rightarrow f)},$$

the  $CP$ -violation charge asymmetry of exclusive  $B^-$  and  $B^+$  decay. $A_{CP}(B^+ \rightarrow J/\psi(1S) K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.003 \pm 0.006</math> OUR AVERAGE</b>			Error includes scale factor of 1.8. See the ideogram below.
$0.0059 \pm 0.0036 \pm 0.0007$	ABAZOV	13M	D0 $p\bar{p}$ at 1.96 TeV
$-0.0076 \pm 0.0050 \pm 0.0022$	SAKAI	10	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.09 \pm 0.07 \pm 0.02$	<sup>1</sup> WEI	08	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.030 \pm 0.014 \pm 0.010$	<sup>2</sup> AUBERT	05J	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.018 \pm 0.043 \pm 0.004$	<sup>3</sup> BONVICINI	00	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.0075 \pm 0.0061 \pm 0.0030$	<sup>4</sup> ABAZOV	08o	D0 Repl. by ABAZOV 13M
$0.03 \pm 0.015 \pm 0.006$	AUBERT	04P	BABR Repl. by AUBERT 05J
$-0.026 \pm 0.022 \pm 0.017$	ABE	03B	BELL Repl. by SAKAI 10
$0.003 \pm 0.030 \pm 0.004$	AUBERT	02f	BABR Repl. by AUBERT 04P

<sup>1</sup> Uses  $B^+ \rightarrow J/\psi K^+$ , where  $J/\psi \rightarrow p\bar{p}$ .<sup>2</sup> The result reported corresponds to  $-A_{CP}$ .<sup>3</sup> A +0.3% correction is applied due to a slightly higher reconstruction efficiency for the positive kaons.<sup>4</sup> Uses  $J/\psi \rightarrow \mu^+ \mu^-$  decay.WEIGHTED AVERAGE  
 $0.003 \pm 0.006$  (Error scaled by 1.8)

# Meson Particle Listings

$B^\pm$

## $A_{CP}(B^+ \rightarrow J/\psi(1S)\pi^+)$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.1 \pm 2.8</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.		
$-4.2 \pm 4.4 \pm 0.9$	ABAZOV	13M D0	$p\bar{p}$ at 1.96 TeV
$0.5 \pm 2.7 \pm 1.1$	<sup>1</sup> AAIJ	12AC LHCb	$pp$ at 7 TeV
$12.3 \pm 8.5 \pm 0.4$	AUBERT	04P BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-2.3 \pm 16.4 \pm 1.5$	ABE	03B BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-9 \pm 8 \pm 3$	<sup>2</sup> ABAZOV	08o D0	Repl. by ABAZOV 13M
$1 \pm 22 \pm 1$	AUBERT	02F BABR	Repl. by AUBERT 04P
<sup>1</sup> Uses $A_{CP}(B^+ \rightarrow J/\psi K^+) = 0.001 \pm 0.007$ to extract production asymmetry.			
<sup>2</sup> Uses $J/\psi \rightarrow \mu^+\mu^-$ decay.			

## $A_{CP}(B^+ \rightarrow J/\psi\rho^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.11 \pm 0.12 \pm 0.08</math></b>	AUBERT	07AC BABR	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow J/\psi K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.048 \pm 0.029 \pm 0.016</math></b>	<sup>1</sup> AUBERT	05J BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> The result reported corresponds to $-A_{CP}$ .			

## $A_{CP}(B^+ \rightarrow \eta_c K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.07</math> OUR AVERAGE</b>	Error includes scale factor of 2.2.		
$0.040 \pm 0.034 \pm 0.004$	<sup>1</sup> AAIJ	14AF LHCb	$pp$ at 7, 8 TeV
$-0.16 \pm 0.08 \pm 0.02$	<sup>1</sup> WEI	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.046 \pm 0.057 \pm 0.007$	<sup>1</sup> AAIJ	13Au LHCb	Repl. by AAIJ 14AF
<sup>1</sup> Uses $B^+ \rightarrow \eta_c K^+$ , where $\eta_c \rightarrow p\bar{p}$ .			

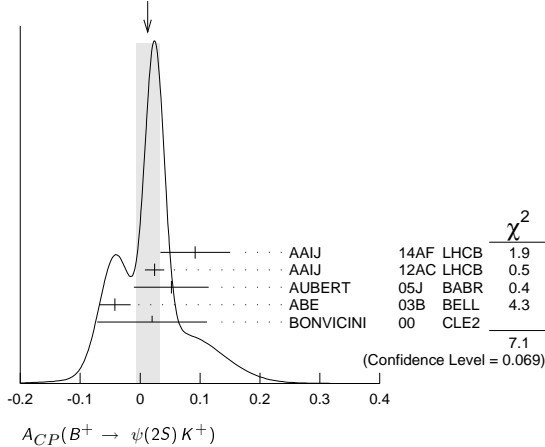
## $A_{CP}(B^+ \rightarrow \psi(2S)\pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.03 \pm 0.06</math> OUR AVERAGE</b>			
$0.048 \pm 0.090 \pm 0.011$	<sup>1</sup> AAIJ	12AC LHCb	$pp$ at 7 TeV
$0.022 \pm 0.085 \pm 0.016$	BHARDWAJ	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses $A_{CP}(B^+ \rightarrow J/\psi K^+) = 0.001 \pm 0.007$ to extract production asymmetry.			

## $A_{CP}(B^+ \rightarrow \psi(2S)K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.012 \pm 0.020</math> OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
$0.092 \pm 0.058 \pm 0.004$	<sup>1</sup> AAIJ	14AF LHCb	$pp$ at 7, 8 TeV
$0.024 \pm 0.014 \pm 0.008$	<sup>2</sup> AAIJ	12AC LHCb	$pp$ at 7 TeV
$0.052 \pm 0.059 \pm 0.020$	AUBERT	05J BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.042 \pm 0.020 \pm 0.017$	ABE	03B BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.02 \pm 0.091 \pm 0.01$	<sup>3</sup> BONVICINI	00 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.002 \pm 0.123 \pm 0.012$	<sup>1,2</sup> AAIJ	13Au LHCb	Repl. by AAIJ 14AF
<sup>1</sup> Uses $\psi(2S) \rightarrow p\bar{p}$ decays.			
<sup>2</sup> Uses $A_{CP}(B^+ \rightarrow J/\psi K^+) = 0.001 \pm 0.007$ to extract production asymmetry.			
<sup>3</sup> A +0.3% correction is applied due to a slightly higher reconstruction efficiency for the positive kaons.			

WEIGHTED AVERAGE  
 $0.012 \pm 0.020$  (Error scaled by 1.5)



## $A_{CP}(B^+ \rightarrow \psi(2S)K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.077 \pm 0.207 \pm 0.051</math></b>	<sup>1</sup> AUBERT	05J BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> The result reported corresponds to $-A_{CP}$ .			

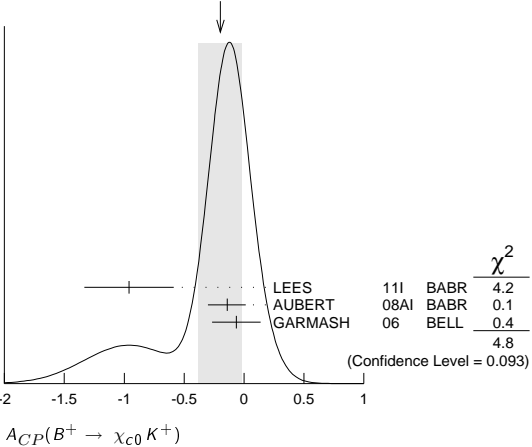
## $A_{CP}(B^+ \rightarrow \chi_{c1}(1P)\pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.07 \pm 0.18 \pm 0.02</math></b>	KUMAR	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \chi_{c0} K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.20 \pm 0.18</math> OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
$-0.96 \pm 0.37 \pm 0.04$	LEES	11i BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.14 \pm 0.15 \pm 0.03$	AUBERT	08AI BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.065 \pm 0.20 \pm 0.035$	GARMASH	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.024$			

WEIGHTED AVERAGE  
 $-0.20 \pm 0.18$  (Error scaled by 1.5)



## $A_{CP}(B^+ \rightarrow \chi_{c1} K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.009 \pm 0.033</math> OUR AVERAGE</b>			
$-0.01 \pm 0.03 \pm 0.02$	KUMAR	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.003 \pm 0.076 \pm 0.017$	<sup>1</sup> AUBERT	05J BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> The result reported corresponds to $-A_{CP}$ .			

## $A_{CP}(B^+ \rightarrow \chi_{c1} K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.471 \pm 0.378 \pm 0.268</math></b>	<sup>1</sup> AUBERT	05J BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> The result reported corresponds to $-A_{CP}$ .			

## $A_{CP}(B^+ \rightarrow \bar{D}^0 \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.007 \pm 0.007</math> OUR AVERAGE</b>			
$-0.006 \pm 0.005 \pm 0.010$	<sup>1</sup> AAIJ	13AE LHCb	$pp$ at 7 TeV
$-0.008 \pm 0.008$	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses $B^\pm \rightarrow [K^\pm \pi^\mp \pi^\pm \pi^-]_D h^\pm$ mode.			

## $A_{CP}(B^+ \rightarrow D_{CP(+1)} \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.035 \pm 0.024</math></b>	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow D_{CP(-1)} \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.017 \pm 0.026</math></b>	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

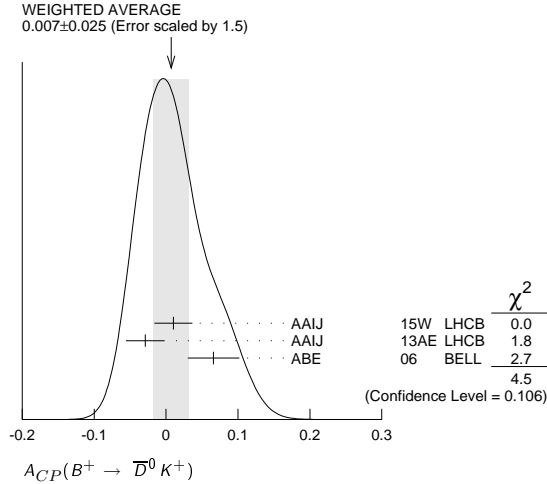
## $A_{CP}([K^\mp \pi^\pm \pi^\pm \pi^-]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.13 \pm 0.10</math></b>	AAIJ	13AE LHCb	$pp$ at 7 TeV

## $A_{CP}(B^+ \rightarrow \bar{D}^0 K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.007 \pm 0.025</math> OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
$0.010 \pm 0.026 \pm 0.005$	<sup>1</sup> AAIJ	15W LHCb	$pp$ at 7, 8 TeV
$-0.029 \pm 0.020 \pm 0.018$	<sup>2</sup> AAIJ	13AE LHCb	$pp$ at 7 TeV
$0.066 \pm 0.036$	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.003 \pm 0.080 \pm 0.037$	<sup>3</sup> ABE	03D BELL	Repl. by SWAIN 03
$0.04 \pm 0.06 \pm 0.03$	<sup>4</sup> SWAIN	03 BELL	Repl. by ABE 06

- <sup>1</sup> Uses  $D^0 \rightarrow K^- \pi^+ \pi^0$  for the favored mode, and  $D^0 \rightarrow K^+ \pi^- \pi^0$  for the suppressed mode.  
<sup>2</sup> Uses  $B^\pm \rightarrow [K^\pm \pi^\mp \pi^+ \pi^-] D h^\pm$  mode.  
<sup>3</sup> Corresponds to 90% confidence range  $-0.15 < A_{CP} < 0.16$ .  
<sup>4</sup> Corresponds to 90% confidence range  $-0.07 < A_{CP} < 0.15$ .



### $A_{CP}([K^\mp \pi^\pm \pi^+ \pi^-] D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.42 \pm 0.22$	AAIJ	13AE LHCB	$pp$ at 7 TeV

### $r_B(B^+ \rightarrow D^0 K^+)$

$r_B$  and  $\delta_B$  are the amplitude ratio and relative strong phase between the amplitudes of  $A(B^+ \rightarrow D^0 K^+)$  and  $A(B^+ \rightarrow \bar{D}^0 K^+)$ ,

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.095 \pm 0.008</math> OUR AVERAGE</b>				
$0.080^{+0.019}_{-0.021}$		<sup>1</sup> AAIJ	14BA LHCB	$pp$ at 7, 8 TeV
$0.07 \pm 0.04$		<sup>6,7</sup> AAIJ	12AQ LHCB	$pp$ at 7 TeV
$0.097 \pm 0.011$		<sup>2</sup> AAIJ	13AE LHCB	$pp$ at 7 TeV
$0.092^{+0.013}_{-0.012}$		<sup>3</sup> LEES	13B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.160^{+0.040+0.051}_{-0.038-0.015}$		<sup>4</sup> POLUEKTOV	10 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.06 \pm 0.04$		<sup>5</sup> AAIJ	14BE LHCB	Repl. by AAIJ 14BA
$0.145 \pm 0.030 \pm 0.015$		<sup>7,8</sup> AIHARA	12 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 0.13$	90	<sup>9</sup> LEES	11D BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.096 \pm 0.029 \pm 0.006$		<sup>10</sup> DEL-AMO-SA..10F	BABR	Repl. by LEES 13B
$0.095^{+0.051}_{-0.041}$		<sup>11</sup> DEL-AMO-SA..10H	BABR	Repl. by LEES 13B
$0.086 \pm 0.032 \pm 0.015$		<sup>12</sup> AUBERT	08AL BABR	Repl. by DEL-AMO-SANCHEZ 10F
$< 0.19$	90	HORI	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.159^{+0.054}_{-0.050} \pm 0.050$		<sup>13</sup> POLUEKTOV	06 BELL	Repl. by POLUEKTOV 10
$0.12 \pm 0.08 \pm 0.05$		<sup>14</sup> AUBERT,B	05Y BABR	Repl. by AUBERT 08AL

- • • We do not use the following data for averages, fits, limits, etc. • • •
- <sup>1</sup> Uses binned Dalitz plot analysis of  $B^+ \rightarrow DK^+$  decays, with  $D \rightarrow K_S^0 \pi^+ \pi^-$  and  $D \rightarrow K_S^0 K^+ K^-$ . Strong phase measurements from CLEO-c (LIBBY 10) of the  $D$  decay over the Dalitz plot are used as input.  
<sup>2</sup> Uses  $B^\pm \rightarrow [K^\pm \pi^\mp \pi^+ \pi^-] D h^\pm$  mode.  
<sup>3</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods.  
<sup>4</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow D^0 K^+$  modes. The corresponding two standard deviation interval is  $0.084 < r_B < 0.239$ .  
<sup>5</sup> AAIJ 14BE uses model-dependent analysis of  $D \rightarrow K_S^0 \pi^+ \pi^-$  amplitudes. The model is the same as in DEL-AMO-SANCHEZ 10F.  
<sup>6</sup> Reports combined statistical and systematic uncertainties.  
<sup>7</sup> Uses binned Dalitz plot of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow \bar{D}^0 K^+$ . Measurement of strong phases in  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  Dalitz plot from LIBBY 10 is used as input.  
<sup>8</sup> We combined the systematics in quadrature. The authors report separately the contribution to the systematic uncertainty due to the uncertainty on the bin-averaged strong phase difference between  $D^0$  and  $\bar{D}^0$  amplitudes.  
<sup>9</sup> Uses decays of neutral  $D$  to  $K^- \pi^+ \pi^0$ .  
<sup>10</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$ ,  $K_S^0 K^+ K^-$  decays from  $B^+ \rightarrow D(*) K(*)^+$  modes. The corresponding two standard deviation interval is  $0.037 < r_B < 0.155$ .  
<sup>11</sup> Uses the Cabibbo suppressed decay of  $B^+ \rightarrow \bar{D} K^+$  followed by  $\bar{D} \rightarrow K^- \pi^+$ .  
<sup>12</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\bar{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D(*) K(*)^\pm$  modes.  
<sup>13</sup> Uses a Dalitz plot analysis of the  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $DK^+$ ,  $D^* K^+$  and  $DK^{*+}$  modes.  
<sup>14</sup> Uses a Dalitz analysis of neutral  $D$  decays to  $K_S^0 \pi^+ \pi^-$  in the processes  $B^\pm \rightarrow D(*) K^\pm$ ,  $D^* \rightarrow D \pi^0$ ,  $D \gamma$ .

### $\delta_B(B^+ \rightarrow D^0 K^+)$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>123 \pm 10</math> OUR AVERAGE</b>			
$134^{+14}_{-15}$	<sup>1</sup> AAIJ	14BA LHCB	$pp$ at 7, 8 TeV
$105^{+16}_{-17}$	<sup>2</sup> LEES	13B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$136.7^{+13.0}_{-15.8} \pm 23.2$	<sup>3</sup> POLUEKTOV	10 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$115^{+41}_{-51}$	<sup>4</sup> AAIJ	14BE LHCB	Repl. by AAIJ 14BA
$137^{+35}_{-46}$	<sup>5,6</sup> AAIJ	12AQ LHCB	$pp$ at 7 TeV
$129.9 \pm 15.0 \pm 6.0$	<sup>6,7</sup> AIHARA	12 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$119^{+19}_{-20} \pm 4$	<sup>8</sup> DEL-AMO-SA..10F	BABR	Repl. by LEES 13B
$109^{+27}_{-30} \pm 8$	<sup>9</sup> AUBERT	08AL BABR	Repl. by DEL-AMO-SANCHEZ 10F
$145.7^{+19.0}_{-19.7} \pm 23.1$	<sup>10</sup> POLUEKTOV	06 BELL	Repl. by POLUEKTOV 10
$104 \pm 45^{+23}_{-32}$	<sup>11</sup> AUBERT,B	05Y BABR	Repl. by AUBERT 08AL

- <sup>1</sup> Uses binned Dalitz plot analysis of  $B^+ \rightarrow DK^+$  decays, with  $D \rightarrow K_S^0 \pi^+ \pi^-$  and  $D \rightarrow K_S^0 K^+ K^-$ . Strong phase measurements from CLEO-c (LIBBY 10) of the  $D$  decay over the Dalitz plot are used as input.  
<sup>2</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods.  
<sup>3</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow \bar{D}^0 K^+$  modes. The corresponding two standard deviation interval is  $102.2^\circ < \delta_B < 162.3^\circ$ .  
<sup>4</sup> AAIJ 14BE uses model-dependent analysis of  $D \rightarrow K_S^0 \pi^+ \pi^-$  amplitudes. The model is the same as in DEL-AMO-SANCHEZ 10F.  
<sup>5</sup> Reports combined statistical and systematic uncertainties.  
<sup>6</sup> Uses binned Dalitz plot of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow \bar{D}^0 K^+$ . Measurement of strong phases in  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  Dalitz plot from LIBBY 10 is used as input.  
<sup>7</sup> We combined the systematics in quadrature. The authors report separately the contribution to the systematic uncertainty due to the uncertainty on the bin-averaged strong phase difference between  $D^0$  and  $\bar{D}^0$  amplitudes.  
<sup>8</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$ ,  $K_S^0 K^+ K^-$  decays from  $B^+ \rightarrow D(*) K(*)^+$  modes. The corresponding two standard deviation interval is  $75^\circ < \delta_B < 157^\circ$ .  
<sup>9</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\bar{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D(*) K(*)^\pm$  modes.  
<sup>10</sup> Uses a Dalitz plot analysis of the  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $DK^+$ ,  $D^* K^+$  and  $DK^{*+}$  modes.  
<sup>11</sup> Uses a Dalitz analysis of neutral  $D$  decays to  $K_S^0 \pi^+ \pi^-$  in the processes  $B^\pm \rightarrow D(*) K^\pm$ ,  $D^* \rightarrow D \pi^0$ ,  $D \gamma$ .

### $r_B(B^+ \rightarrow \bar{D}^0 K^{*+})$

$r_B$  and  $\delta_B$  are the amplitude ratio and relative strong phase between the amplitudes of  $A_{CP}(B^+ \rightarrow D^0 K^{*+})$  and  $A_{CP}(B^+ \rightarrow \bar{D}^0 K^{*+})$ ,

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.17 \pm 0.11</math> OUR AVERAGE</b>			Error includes scale factor of 2.3.
$0.143^{+0.048}_{-0.049}$	<sup>1</sup> LEES	13B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.564^{+0.216}_{-0.155} \pm 0.093$	<sup>2</sup> POLUEKTOV	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.166^{+0.073}_{-0.069}$	<sup>3</sup> DEL-AMO-SA..10F	BABR	Repl. by LEES 13B
$0.31 \pm 0.07$	<sup>4</sup> AUBERT	09AJ BABR	Repl. by LEES 13B
$0.181^{+0.088}_{-0.108} \pm 0.042$	<sup>5</sup> AUBERT	08AL BABR	Repl. by AUBERT 09AJ

- <sup>1</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods.  
<sup>2</sup> Uses a Dalitz plot analysis of the  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $DK^+$ ,  $D^* K^+$  and  $DK^{*+}$  modes.  
<sup>3</sup> DEL-AMO-SANCHEZ 10F reports  $r_B \cdot k = 0.149^{+0.066}_{-0.062}$  for  $k = 0.9$ .  
<sup>4</sup> Obtained by combining the GLW and ADS methods. The 2-sigma range corresponds to  $[0.17, 0.43]$ .  
<sup>5</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\bar{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D(*) K(*)^\pm$  modes.

### $\delta_B(B^+ \rightarrow D^0 K^{*+})$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>155 \pm 70</math> OUR AVERAGE</b>			Error includes scale factor of 2.0.
$101 \pm 43$	<sup>1</sup> LEES	13B BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$242.6^{+20.2}_{-23.2} \pm 49.4$	<sup>2</sup> POLUEKTOV	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$111 \pm 32$	DEL-AMO-SA..10F	BABR	Repl. by LEES 13B
$104^{+39}_{-37} \pm 18$	<sup>3</sup> AUBERT	08AL BABR	Repl. by LEES 13B

- <sup>1</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods.  
<sup>2</sup> Uses a Dalitz plot analysis of the  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $DK^+$ ,  $D^* K^+$  and  $DK^{*+}$  modes.  
<sup>3</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\bar{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D(*) K(*)^\pm$  modes.

# Meson Particle Listings

$B^\pm$

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.58 \pm 0.21</math> OUR AVERAGE</b>			
$-0.82 \pm 0.44 \pm 0.09$	AALTONEN	11AJ	CDF $p\overline{p}$ at 1.96 TeV
$-0.39 \pm 0.26 \pm 0.04$ $-0.28 \pm 0.03$	HORII	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.86 \pm 0.47 \pm 0.12$ $-0.16$	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.1 \pm 0.8$ $-1.0 \pm 0.4$	HORII	08	BELL Repl. by HORII 11
$+0.88 \pm 0.77 \pm 0.06$ $-0.62$	SAIGO	05	BELL Repl. by HORII 08

## $A_{CP}(B^+ \rightarrow [K^- \pi^+ \pi^0]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.07 \pm 0.30</math> OUR AVERAGE</b>			Error includes scale factor of 1.5.
$-0.20 \pm 0.27 \pm 0.04$	<sup>1</sup> AAIJ	15W	LHCB $pp$ at 7, 8 TeV
$0.41 \pm 0.30 \pm 0.05$	NAYAK	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses $D^0 \rightarrow K^- \pi^+ \pi^0$ for the favored mode, and $D^0 \rightarrow K^+ \pi^- \pi^0$ for the suppressed mode.			

## $A_{CP}(B^+ \rightarrow [K^+ K^- \pi^0]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.30 \pm 0.20 \pm 0.02</math></b>	<sup>1</sup> AAIJ	15W	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Uses $D \rightarrow K^+ K^- \pi^0$ mode.			

## $A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.054 \pm 0.091 \pm 0.011</math></b>	<sup>1</sup> AAIJ	15W	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Uses $D \rightarrow \pi^+ \pi^- \pi^0$ mode.			

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_{\overline{D}} K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.34 \pm 0.43 \pm 0.16</math></b>	AUBERT	09AJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.22 \pm 0.61 \pm 0.17$	AUBERT,B	05V	BABR Repl. by AUBERT 09AJ

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.09</math> OUR AVERAGE</b>			
$0.13 \pm 0.25 \pm 0.02$	AALTONEN	11AJ	CDF $p\overline{p}$ at 1.96 TeV
$-0.04 \pm 0.11 \pm 0.02$ $-0.01$	HORII	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.03 \pm 0.17 \pm 0.04$	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.02 \pm 0.15$ $-0.16 \pm 0.04$	HORII	08	BELL Repl. by HORII 11
$+0.30 \pm 0.29 \pm 0.06$ $-0.25$	SAIGO	05	BELL Repl. by HORII 08

## $A_{CP}(B^+ \rightarrow [K^- \pi^+ \pi^0]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.35 \pm 0.16</math> OUR AVERAGE</b>			
$0.438 \pm 0.190 \pm 0.011$	<sup>1</sup> AAIJ	15W	LHCB $pp$ at 7, 8 TeV
$0.16 \pm 0.27 \pm 0.03$ $-0.04$	NAYAK	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses $D^0 \rightarrow K^- \pi^+ \pi^0$ for the favored mode, and $D^0 \rightarrow K^+ \pi^- \pi^0$ for the suppressed mode.			

## $A_{CP}(B^+ \rightarrow [K^+ K^- \pi^0]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.030 \pm 0.040 \pm 0.005</math></b>	<sup>1</sup> AAIJ	15W	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Uses $D \rightarrow K^+ K^-$ mode.			

## $A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.016 \pm 0.020 \pm 0.004</math></b>	<sup>1</sup> AAIJ	15W	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Uses $D \rightarrow \pi^+ \pi^-$ mode.			

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\pi)} \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.09 \pm 0.27 \pm 0.05</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\gamma)} \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.65 \pm 0.55 \pm 0.22</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\pi)} K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.77 \pm 0.35 \pm 0.12</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow [K^- \pi^+]_{(D\gamma)} K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.36 \pm 0.94 \pm 0.25</math></b> <b><math>-0.41</math></b>	DEL-AMO-SA...10H	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow [\pi^+ \pi^- \pi^0]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.02 \pm 0.15 \pm 0.03</math></b>	<sup>1</sup> AUBERT	07BJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.02 \pm 0.16 \pm 0.03$	AUBERT,B	05T	BABR Repl. by AUBERT 07BJ
<sup>1</sup> Uses a Dalitz plot analysis of $D^0 \rightarrow \pi^+ \pi^- \pi^0$ . Also reports the one-sigma regions: $0.06 < r_B < 0.78$ , $-30^\circ < \gamma < 76^\circ$ , and $-27^\circ < \delta < 78^\circ$ .			

## $A_{CP}(B^+ \rightarrow [K_S^0 K^+ \pi^-]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.040 \pm 0.091 \pm 0.018</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The anaysis uses all of $D \rightarrow K_S^0 K \pi$ Dalitz decays.			

## $A_{CP}(B^+ \rightarrow [K_S^0 K^- \pi^+]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.233 \pm 0.129 \pm 0.024</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The anaysis uses all of $D \rightarrow K_S^0 K \pi$ Dalitz decays.			

## $A_{CP}(B^+ \rightarrow [K_S^0 K^- \pi^+]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.052 \pm 0.029 \pm 0.017</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The anaysis uses all of $D \rightarrow K_S^0 K \pi$ Dalitz decays.			

## $A_{CP}(B^+ \rightarrow [K_S^0 K^+ \pi^-]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.025 \pm 0.024 \pm 0.010</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The anaysis uses all of $D \rightarrow K_S^0 K \pi$ Dalitz decays.			

## $A_{CP}(B^+ \rightarrow [K^*(892)^- K^+]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.026 \pm 0.109 \pm 0.029</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The Analysis uses $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$ decays.			

## $A_{CP}(B^+ \rightarrow [K^*(892)^+ K^-]_D K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.336 \pm 0.208 \pm 0.026</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The Analysis uses $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$ decays.			

## $A_{CP}(B^+ \rightarrow [K^*(892)^+ K^-]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.054 \pm 0.043 \pm 0.017</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The Analysis uses $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$ decays.			

## $A_{CP}(B^+ \rightarrow [K^*(892)^- K^+]_D \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.012 \pm 0.028 \pm 0.010</math></b>	<sup>1</sup> AAIJ	14V	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> The Analysis uses $D \rightarrow K^*(892) K \rightarrow K_S^0 K \pi$ decays.			

## $A_{CP}(B^+ \rightarrow D_{CP(+1)} K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.170 \pm 0.033</math> OUR AVERAGE</b>			Error includes scale factor of 1.2.
$0.145 \pm 0.032 \pm 0.010$	<sup>1</sup> AAIJ	12M	LHCB $pp$ at 7 TeV
$0.39 \pm 0.17 \pm 0.04$	AALTONEN	10A	CDF $p\overline{p}$ at 1.96 TeV
$0.25 \pm 0.06 \pm 0.02$	<sup>2</sup> DEL-AMO-SA...10G	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.06 \pm 0.14 \pm 0.05$	ABE	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.27 \pm 0.09 \pm 0.04$	AUBERT	08AA	BABR Repl. by DEL-AMO-SANCHEZ 10G
$0.35 \pm 0.13 \pm 0.04$	AUBERT	06J	BABR Repl. by AUBERT 08AA
$0.07 \pm 0.17 \pm 0.06$	AUBERT	04N	BABR Repl. by AUBERT 06J
$0.29 \pm 0.26 \pm 0.05$	<sup>3</sup> ABE	03D	BELL Repl. by SWAIN 03
$0.06 \pm 0.19 \pm 0.04$	<sup>4</sup> SWAIN	03	BELL Repl. by ABE 06

<sup>1</sup> AAIJ 12M reports an evidence of direct  $CP$  violation in  $B^\pm \rightarrow D K^\pm$  decays with a total significance of 5.8  $\sigma$ .

<sup>2</sup> Reports the first evidence for direct  $CP$  violation in  $B \rightarrow D K$  decays with 3.6 standard deviations.

<sup>3</sup> Corresponds to 90% confidence range  $-0.14 < A_{CP} < 0.73$ .

<sup>4</sup> Corresponds to 90% confidence range  $-0.26 < A_{CP} < 0.38$ .

## $A_{ADS}(B^+ \rightarrow DK^+)$

$$A_{ADS}(B^+ \rightarrow DK^+) = \frac{(R_K^- - R_K^+)}{(R_K^- + R_K^+)} \text{ where}$$

$$R_K^- = \Gamma(B^- \rightarrow [K^+ \pi^-]_D K^-) / \Gamma(B^- \rightarrow [K^- \pi^+]_D K^-) \text{ and}$$

$$R_K^+ = \Gamma(B^+ \rightarrow [K^- \pi^+]_D K^+) / \Gamma(B^+ \rightarrow [K^+ \pi^-]_D K^+)$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.52 \pm 0.15 \pm 0.02</math></b>	AAIJ	12M	LHCB $pp$ at 7 TeV

$A_{ADS}(B^+ \rightarrow D\pi^+)$ 

$$A_{ADS}(B^+ \rightarrow D\pi^+) = \frac{(R_\pi^- - R_\pi^+)}{(R_\pi^- + R_\pi^+)} \text{ where}$$

$$R_\pi^- = \Gamma(B^- \rightarrow [K^+\pi^-]_D \pi^-) / \Gamma(B^- \rightarrow [K^-\pi^+]_D \pi^-) \text{ and}$$

$$R_\pi^+ = \Gamma(B^+ \rightarrow [K^-\pi^+]_D \pi^+) / \Gamma(B^+ \rightarrow [K^+\pi^-]_D \pi^+)$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.143 \pm 0.062 \pm 0.011</math></b>	AAIJ	12M LHCb	$pp$ at 7 TeV

 $A_{ADS}(B^+ \rightarrow [K^-\pi^+]_D K^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.33^{+0.36}_{-0.34}</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{ADS}(B^+ \rightarrow [K^-\pi^+]_D \pi^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.013 \pm 0.087</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow D_{CP(-1)} K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.10 \pm 0.07</math> OUR AVERAGE</b>			
$-0.09 \pm 0.07 \pm 0.02$	DEL-AMO-SA...10g	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.12 \pm 0.14 \pm 0.05$	ABE 06	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.09 \pm 0.09 \pm 0.02$	AUBERT	08AA BABR	Repl. by DEL-AMO-SANCHEZ 10G
$-0.06 \pm 0.13 \pm 0.04$	AUBERT	06J BABR	Repl. by AUBERT 08AA
$-0.22 \pm 0.14 \pm 0.04$	<sup>1</sup> ABE	03D BELL	Repl. by SWAIN 03
$-0.19 \pm 0.17 \pm 0.05$	<sup>2</sup> SWAIN	03 BELL	Repl. by ABE 06

<sup>1</sup> Corresponds to 90% confidence range  $-0.62 < A_{CP} < 0.18$ .

<sup>2</sup> Corresponds to 90% confidence range  $-0.47 < A_{CP} < 0.11$ .

 $A_{CP}(B^+ \rightarrow [K^+K^-]_D K^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.045 \pm 0.064 \pm 0.011</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow [\pi^+\pi^-]_D K^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.054 \pm 0.101 \pm 0.011</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow [K^-\pi^+]_D K^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.013 \pm 0.019 \pm 0.013</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow [K^+K^-]_D \pi^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.019 \pm 0.011 \pm 0.010</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow [\pi^+\pi^-]_D \pi^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.013 \pm 0.016 \pm 0.010</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow [K^-\pi^+]_D \pi^+\pi^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.002 \pm 0.003 \pm 0.011</math></b>	AAIJ	15bc LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow \bar{D}^{*0} \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.014 \pm 0.015</math></b>	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow (D_{CP(+1)}^*)^0 \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.021 \pm 0.045</math></b>	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow (D_{CP(-1)}^*)^0 \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.090 \pm 0.051</math></b>	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow D^{*0} K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.07 \pm 0.04</math> OUR AVERAGE</b>			
$-0.06 \pm 0.04 \pm 0.01$	AUBERT	08Bf BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.089 \pm 0.086$	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $r_B^*(B^+ \rightarrow D^{*0} K^+)$ 

$r_B^*$  and  $\delta_B^*$  are the amplitude ratio and relative strong phase between the amplitudes of  $A(B^+ \rightarrow D^{*0} K^+)$  and  $A(B^+ \rightarrow \bar{D}^{*0} K^+)$ ,

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.114^{+0.023}_{-0.040}</math> OUR AVERAGE</b>			Error includes scale factor of 1.2.
$0.106^{+0.019}_{-0.036}$	<sup>1</sup> LEES	13B BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.196^{+0.072+0.064}_{-0.069-0.017}$	<sup>2</sup> POLUEKTOV	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.133^{+0.042}_{-0.039} \pm 0.013$	<sup>3</sup> DEL-AMO-SA...10F	BABR	Repl. by LEES 13B
$0.096^{+0.035}_{-0.051}$	<sup>4</sup> DEL-AMO-SA...10H	BABR	Repl. by LEES 13B
$0.135 \pm 0.050 \pm 0.012$	<sup>5</sup> AUBERT	08AL BABR	Repl. by DEL-AMO-SANCHEZ 10F
$0.175^{+0.108}_{-0.099} \pm 0.050$	<sup>6</sup> POLUEKTOV	06 BELL	Repl. by POLUEKTOV 10
$0.17 \pm 0.10 \pm 0.04$	<sup>7</sup> AUBERT,B	05Y BABR	Repl. by AUBERT 08AL

<sup>1</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods.

<sup>2</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow D^{*0} K^+$  modes. The corresponding two standard deviation interval is  $0.061 < r_B^* < 0.271$ .

<sup>3</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$ ,  $K_S^0 K^+ K^-$  decays from  $B^+ \rightarrow D^{(*)} K^{(*)+}$  modes. The corresponding two standard deviation interval is  $0.049 < r_B^* < 0.215$ .

<sup>4</sup> Uses the Cabibbo suppressed decay of  $B^+ \rightarrow \bar{D}^* K^+$  followed by  $\bar{D}^* \rightarrow \bar{D} \pi^0$  or  $\bar{D} \gamma$ , and  $\bar{D} \rightarrow K^-\pi^+$ .

<sup>5</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\bar{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D^{(*)} K^{(*)\pm}$  modes.

<sup>6</sup> Uses a Dalitz plot analysis of the  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $D K^+$ ,  $D^* K^+$  and  $D K^{*+}$  modes.

<sup>7</sup> Uses a Dalitz plot analysis of neutral  $D$  decays to  $K_S^0 \pi^+ \pi^-$  in the processes  $B^\pm \rightarrow D^{(*)} K^\pm$ ,  $D^* \rightarrow D \pi^0$ ,  $D \gamma$ .

 $\delta_B^*(B^+ \rightarrow D^{*0} K^+)$ 

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>310^{+22}_{-28}</math> OUR AVERAGE</b>			Error includes scale factor of 1.3.

294 $^{+21}_{-31}$	<sup>1</sup> LEES	13B BABR	$e^+e^- \rightarrow \Upsilon(4S)$
341.9 $^{+18.0}_{-19.6} \pm 23.1$	<sup>2</sup> POLUEKTOV	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

278 $\pm 21 \pm 6$	<sup>3</sup> DEL-AMO-SA...10F	BABR	Repl. by LEES 13B
297 $^{+27}_{-29} \pm 6.4$	<sup>4</sup> AUBERT	08AL BABR	Repl. by DEL-AMO-SANCHEZ 10F
302.0 $^{+33.8}_{-35.1} \pm 23.7$	<sup>5</sup> POLUEKTOV	06 BELL	Repl. by POLUEKTOV 10
296 $\pm 41^{+20}_{-19}$	<sup>6</sup> AUBERT,B	05Y BABR	Repl. by AUBERT 08AL

<sup>1</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods. We added  $360^\circ$  to the value of  $(-66^{+21}_{-31})^\circ$  quoted by LEES 13B.

<sup>2</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow D^* K^+$  modes. The corresponding two standard deviation interval is  $296.5^\circ < \delta_B^* < 382.7^\circ$ .

<sup>3</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$ ,  $K_S^0 K^+ K^-$  decays from  $B^+ \rightarrow D^{(*)} K^{(*)+}$  modes. The corresponding two standard deviation interval is  $236^\circ < \delta_B^* < 322^\circ$ .

<sup>4</sup> Uses Dalitz plot analysis of  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\bar{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D^{(*)} K^{(*)\pm}$  modes.

<sup>5</sup> Uses a Dalitz plot analysis of the  $\bar{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $D K^+$ ,  $D^* K^+$  and  $D K^{*+}$  modes.

<sup>6</sup> Uses a Dalitz plot analysis of neutral  $D$  decays to  $K_S^0 \pi^+ \pi^-$  in the processes  $B^\pm \rightarrow D^{(*)} K^\pm$ ,  $D^* \rightarrow D \pi^0$ ,  $D \gamma$ .

 $A_{CP}(B^+ \rightarrow D_{CP(+1)}^{*0} K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.12 \pm 0.08</math> OUR AVERAGE</b>			
$-0.11 \pm 0.09 \pm 0.01$	AUBERT	08Bf BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.20 \pm 0.22 \pm 0.04$	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.10 \pm 0.23^{+0.03}_{-0.04}$	AUBERT	05N BABR	Repl. by AUBERT 08Bf

 $A_{CP}(B^+ \rightarrow D_{CP(-1)}^{*0} K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.07 \pm 0.10</math> OUR AVERAGE</b>			
$+0.06 \pm 0.10 \pm 0.02$	AUBERT	08Bf BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$+0.13 \pm 0.30 \pm 0.08$	ABE	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow D_{CP(+1)} K^*(892)^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>+0.09 \pm 0.13 \pm 0.06</math></b>	AUBERT	09AJ BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.08 \pm 0.19 \pm 0.08$	AUBERT,B	05U BABR	Repl. by AUBERT 09AJ

 $A_{CP}(B^+ \rightarrow D_{CP(-1)} K^*(892)^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.23 \pm 0.21 \pm 0.07</math></b>	AUBERT	09AJ BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.26 \pm 0.40 \pm 0.12$	AUBERT,B	05U BABR	Repl. by AUBERT 09AJ

# Meson Particle Listings

$B^\pm$

## $A_{CP}(B^+ \rightarrow D_s^+ \phi)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.01 \pm 0.41 \pm 0.03$	AAIJ	13R	LHCB $pp$ at 7 TeV

## $A_{CP}(B^+ \rightarrow D^{*+} \bar{D}^{*0})$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.15 \pm 0.11 \pm 0.02$	AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow D^{*+} \bar{D}^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.06 \pm 0.13 \pm 0.02$	AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow D^+ \bar{D}^{*0})$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.13 \pm 0.18 \pm 0.04$	AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow D^+ \bar{D}^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.03 \pm 0.07$ OUR AVERAGE			
$0.00 \pm 0.08 \pm 0.02$	ADACHI	08	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.13 \pm 0.14 \pm 0.02$	AUBERT,B	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow K_S^0 \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.017 \pm 0.016$ OUR AVERAGE			
$-0.022 \pm 0.025 \pm 0.010$	AAIJ	13Bs	LHCB $pp$ at 7 TeV
$-0.011 \pm 0.021 \pm 0.006$	DUH	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.029 \pm 0.039 \pm 0.010$	<sup>1</sup> AUBERT,BE	06c	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.18 \pm 0.24$	<sup>2</sup> CHEN	00	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.03 \pm 0.03 \pm 0.01$	LIN	07	BELL Repl. by DUH 13
$-0.09 \pm 0.05 \pm 0.01$	<sup>3</sup> AUBERT,BE	05E	BABR Repl. by AUBERT,BE 06c
$0.05 \pm 0.05 \pm 0.01$	<sup>4</sup> CHAO	05A	BELL Repl. by LIN 07
$-0.05 \pm 0.08 \pm 0.01$	<sup>5</sup> AUBERT	04M	BABR Repl. by AUBERT,BE 05E
$0.07 \pm 0.09 \pm 0.01$	<sup>6</sup> UNNO	03	BELL Repl. by CHAO 05A
$-0.08 \pm 0.03$	<sup>7</sup> CASEY	02	BELL Repl. by UNNO 03
$0.46 \pm 0.15 \pm 0.02$	<sup>8</sup> ABE	01K	BELL Repl. by CASEY 02
$0.098 \pm 0.430 \pm 0.020$	<sup>9</sup> AUBERT	01E	BABR Repl. by AUBERT 04M
$-0.343 \pm 0.063$			
$-0.21 \pm 0.18 \pm 0.03$			

- <sup>1</sup> Corresponds to 90% confidence range  $-0.092 < A_{CP} < 0.036$ .  
<sup>2</sup> Corresponds to 90% confidence range  $-0.22 < A_{CP} < 0.56$ .  
<sup>3</sup> Corresponds to 90% confidence range  $-0.16 < A_{CP} < -0.02$ .  
<sup>4</sup> Corresponds to 90% confidence range  $-0.04 < A_{CP} < 0.13$ .  
<sup>5</sup> Corresponds to 90% confidence range  $-0.18 < A_{CP} < 0.08$ .  
<sup>6</sup> Corresponds to 90% confidence range  $-0.10 < A_{CP} < +0.22$ .  
<sup>7</sup> Corresponds to 90% confidence range  $+0.19 < A_{CP} < +0.72$ .  
<sup>8</sup> Corresponds to 90% confidence range  $-0.53 < A_{CP} < 0.82$ .  
<sup>9</sup> Corresponds to 90% confidence range  $-0.51 < A_{CP} < 0.09$ .

## $A_{CP}(B^+ \rightarrow K^+ \pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.037 \pm 0.021$ OUR AVERAGE			
$0.043 \pm 0.024 \pm 0.002$	DUH	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.030 \pm 0.039 \pm 0.010$	AUBERT	07Bc	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.29 \pm 0.23$	<sup>1</sup> CHEN	00	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.07 \pm 0.03 \pm 0.01$	LIN	08	BELL Repl. by DUH 13
$0.06 \pm 0.06 \pm 0.01$	<sup>2</sup> AUBERT	05L	BABR Repl. by AUBERT 07Bc
$0.06 \pm 0.06 \pm 0.02$	<sup>2</sup> CHAO	05A	BELL Repl. by CHAO 04B
$0.04 \pm 0.05 \pm 0.02$	<sup>3</sup> CHAO	04B	BELL Repl. by LIN 08
$-0.09 \pm 0.09 \pm 0.01$	<sup>4</sup> AUBERT	03L	BABR Repl. by AUBERT 05L
$-0.02 \pm 0.19 \pm 0.02$	<sup>5</sup> CASEY	02	BELL Repl. by CHAO 04B
$-0.059 \pm 0.222 \pm 0.055$	<sup>6</sup> ABE	01K	BELL Repl. by CASEY 02
$-0.196 \pm 0.017$	<sup>7</sup> AUBERT	01E	BABR Repl. by AUBERT 03L
$0.00 \pm 0.18 \pm 0.04$			

- <sup>1</sup> Corresponds to 90% confidence range  $-0.67 < A_{CP} < 0.09$ .  
<sup>2</sup> Corresponds to a 90% CL interval of  $-0.06 < A_{CP} < 0.18$ .  
<sup>3</sup> Corresponds to 90% CL interval of  $-0.05 < A_{CP} < 0.13$ .  
<sup>4</sup> Corresponds to 90% confidence range  $-0.24 < A_{CP} < 0.06$ .  
<sup>5</sup> Corresponds to 90% confidence range  $-0.35 < A_{CP} < +0.30$ .  
<sup>6</sup> Corresponds to 90% confidence range  $-0.40 < A_{CP} < 0.36$ .  
<sup>7</sup> Corresponds to 90% confidence range  $-0.30 < A_{CP} < +0.30$ .

## $A_{CP}(B^+ \rightarrow \eta' K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.004 \pm 0.011$ OUR AVERAGE			
$-0.002 \pm 0.012 \pm 0.006$	<sup>1</sup> AAIJ	15o	LHCB $pp$ at 7, 8 TeV
$0.008 \pm 0.017 \pm 0.009$	AUBERT	09Av	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.028 \pm 0.028 \pm 0.021$	SCHUEMANN	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.03 \pm 0.12$	<sup>2</sup> CHEN	00	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.010 \pm 0.022 \pm 0.006$	AUBERT	07Ae	BABR Repl. by AUBERT 09Av
$0.033 \pm 0.028 \pm 0.005$	<sup>3</sup> AUBERT	05M	BABR Repl. by AUBERT 07Ae
$0.037 \pm 0.045 \pm 0.011$	<sup>4</sup> AUBERT	03W	BABR Repl. by AUBERT 05M
$-0.11 \pm 0.11 \pm 0.02$	<sup>5</sup> AUBERT	02E	BABR Repl. by AUBERT 05M
$-0.015 \pm 0.070 \pm 0.009$	<sup>6</sup> CHEN	02b	BELL Repl. by SCHUEMANN 06
$0.06 \pm 0.15 \pm 0.01$	<sup>7</sup> ABE	01M	BELL Repl. by CHEN 02b
<sup>1</sup> Obtained using $A_{CP}(B^\pm \rightarrow J/\psi K^\pm) = (0.3 \pm 0.6) \times 10^{-2}$ .			
<sup>2</sup> Corresponds to 90% confidence range $-0.17 < A_{CP} < 0.23$ .			
<sup>3</sup> Corresponds to 90% confidence range $-0.012 < A_{CP} < 0.078$ .			
<sup>4</sup> Corresponds to 90% confidence range $-0.04 < A_{CP} < 0.11$ .			
<sup>5</sup> Corresponds to 90% confidence range $-0.28 < A_{CP} < 0.07$ .			
<sup>6</sup> Corresponds to 90% confidence range $-0.13 < A_{CP} < 0.10$ .			
<sup>7</sup> Corresponds to 90% confidence range $-0.20 < A_{CP} < 0.32$ .			

## $A_{CP}(B^+ \rightarrow \eta' K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.26 \pm 0.27 \pm 0.02$	DEL-AMO-SA...10A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.30 \pm 0.33 \pm 0.02$	<sup>1</sup> AUBERT	07E	BABR Repl. by DEL-AMO-SANCHEZ 10A
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<sup>1</sup> Reports  $A_{CP}$  with the opposite sign convention.

## $A_{CP}(B^+ \rightarrow \eta' K_0^*(1430)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.06 \pm 0.20 \pm 0.02$	DEL-AMO-SA...10A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \eta' K_2^*(1430)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.15 \pm 0.13 \pm 0.02$	DEL-AMO-SA...10A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \eta K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.37 \pm 0.08$ OUR AVERAGE			
$-0.38 \pm 0.11 \pm 0.01$	HOI	12	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.36 \pm 0.11 \pm 0.03$	AUBERT	09Av	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.22 \pm 0.11 \pm 0.01$	AUBERT	07Ae	BABR Repl. by AUBERT 09Av
$-0.39 \pm 0.16 \pm 0.03$	CHANG	07B	BELL Repl. by HOI 12
$-0.20 \pm 0.15 \pm 0.01$	AUBERT,B	05K	BABR Repl. by AUBERT 07Ae
$-0.49 \pm 0.31 \pm 0.07$	CHANG	05A	BELL Repl. by CHANG 07B
$-0.52 \pm 0.24 \pm 0.01$	AUBERT	04H	BABR Repl. by AUBERT,B 05K

## $A_{CP}(B^+ \rightarrow \eta K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.02 \pm 0.06$ OUR AVERAGE			
$0.03 \pm 0.10 \pm 0.01$	WANG	07B	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.01 \pm 0.08 \pm 0.02$	AUBERT,B	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.13 \pm 0.14 \pm 0.02$	AUBERT,B	04D	BABR Repl. by AUBERT,B 06H

## $A_{CP}(B^+ \rightarrow \eta K_0^*(1430)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.05 \pm 0.13 \pm 0.02$	AUBERT,B	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \eta K_2^*(1430)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.45 \pm 0.30 \pm 0.02$	AUBERT,B	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \omega K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.02 \pm 0.04$ OUR AVERAGE			
$-0.03 \pm 0.04 \pm 0.01$	CHOBANOVA	14	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.01 \pm 0.07 \pm 0.01$	AUBERT	07Ae	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.05 \pm 0.09 \pm 0.01$	AUBERT,B	06E	BABR Repl. by AUBERT 07Ae
$0.05 \pm 0.08 \pm 0.01$	JEN	06	BELL Repl. by CHOBANOVA 14
$-0.09 \pm 0.17 \pm 0.01$	AUBERT	04H	BABR Repl. by AUBERT,B 06E
$0.06 \pm 0.21 \pm 0.18 \pm 0.01$	<sup>1</sup> WANG	04A	BELL Repl. by JEN 06
$-0.21 \pm 0.28 \pm 0.03$	<sup>2</sup> LU	02	BELL Repl. by WANG 04A
<sup>1</sup> Corresponds to 90% CL interval $0.15 < A_{CP} < 0.90$			
<sup>2</sup> Corresponds to 90% confidence range $-0.70 < A_{CP} < +0.38$ .			

## $A_{CP}(B^+ \rightarrow \omega K^{*+})$

VALUE	DOCUMENT ID	TECN	COMMENT
$+0.29 \pm 0.35 \pm 0.02$	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \omega(K\pi)_0^{*+})$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.10 \pm 0.09 \pm 0.02$	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^+ \rightarrow \omega K_2^*(1430)^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$+0.14 \pm 0.15 \pm 0.02$	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^{*0} \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.04 \pm 0.09$ OUR AVERAGE	Error includes scale factor of 2.1.		
$0.032 \pm 0.052^{+0.016}_{-0.013}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.149 \pm 0.064 \pm 0.022$	GARMASH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.068 \pm 0.078^{+0.070}_{-0.067}$	AUBERT,B	05N	BABR Repl. by AUBERT 08Ai

 $A_{CP}(B^+ \rightarrow K^*(892)^+ \pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.06 \pm 0.24 \pm 0.04$	LEES	11i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.04 \pm 0.29 \pm 0.05$	AUBERT	05x	BABR Repl. by LEES 11i

 $A_{CP}(B^+ \rightarrow K^+ \pi^- \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.027 \pm 0.008$ OUR AVERAGE			
$0.025 \pm 0.004 \pm 0.008$	1 AAIJ	14Bo	LHCB $pp$ at 7, 8 TeV
$0.028 \pm 0.020 \pm 0.023$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.049 \pm 0.026 \pm 0.020$	GARMASH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.032 \pm 0.008 \pm 0.008$	AAIJ	13AZ	LHCB Repl. by AAIJ 14Bo
$-0.013 \pm 0.037 \pm 0.011$	AUBERT,B	05N	BABR Repl. by AUBERT 08Ai
$0.01 \pm 0.07 \pm 0.03$	AUBERT	03M	BABR Repl. by AUBERT,B 05N
1 AAIJ 14Bo reports also $CP$ asymmetries in restricted regions of phase space.			

 $A_{CP}(B^+ \rightarrow K^+ K^- K^+ \text{nonresonant})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.060 \pm 0.044 \pm 0.019$	LEES	12o	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow f(980)^0 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08 \pm 0.08 \pm 0.04$	1 LEES	12o	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
1 Measured in the $B^+ \rightarrow K^+ K^- K^+$ decay.			

 $A_{CP}(B^+ \rightarrow f_2(1270) K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.68^{+0.19}_{-0.17}$ OUR AVERAGE			
$-0.85 \pm 0.22^{+0.26}_{-0.13}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.59 \pm 0.22 \pm 0.036$	GARMASH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow f_0(1500) K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.28 \pm 0.26^{+0.15}_{-0.14}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow f_2'(1525)^0 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08^{+0.05}_{-0.04}$ OUR AVERAGE			
$0.18 \pm 0.18 \pm 0.04$	1 LEES	11i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.106 \pm 0.050^{+0.036}_{-0.015}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.077 \pm 0.065^{+0.046}_{-0.026}$	GARMASH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.14 \pm 0.10 \pm 0.04$	2 LEES	12o	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.31 \pm 0.25 \pm 0.08$	3 AUBERT	06o	BABR Repl. by LEES 12o
$0.088 \pm 0.095^{+0.097}_{-0.056}$	AUBERT,B	05N	BABR Repl. by AUBERT 08Ai
1 Measured in $B^+ \rightarrow f_0 K^+$ with $f_0 \rightarrow \pi^0 \pi^0$ decay.			
2 Measured in the $B^+ \rightarrow K^+ K^- K^+$ decay assuming $A_{CP}(B^+ \rightarrow f_2'(1525)^0 K^+) = A_{CP}(B^+ \rightarrow f_0(1500)^0 K^+) = A_{CP}(B^+ \rightarrow f_0(1710)^0 K^+)$			
3 Measured in the $B^+ \rightarrow K^+ K^- K^+$ decay.			

 $A_{CP}(B^+ \rightarrow \rho^0 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.37 \pm 0.10$ OUR AVERAGE			
$0.44 \pm 0.10^{+0.06}_{-0.14}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.30 \pm 0.11^{+0.11}_{-0.04}$	GARMASH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.32 \pm 0.13^{+0.10}_{-0.08}$	AUBERT,B	05N	BABR Repl. by AUBERT 08Ai

 $A_{CP}(B^+ \rightarrow K_0^*(1430)^0 \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.055 \pm 0.033$ OUR AVERAGE			
$0.032 \pm 0.035^{+0.034}_{-0.028}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.076 \pm 0.038^{+0.028}_{-0.022}$	GARMASH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.064 \pm 0.032^{+0.023}_{-0.026}$	AUBERT,B	05N	BABR Repl. by AUBERT 08Ai

 $A_{CP}(B^+ \rightarrow K_2^*(1430)^0 \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.05 \pm 0.23^{+0.18}_{-0.08}$	AUBERT	08Ai	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^+ \pi^0 \pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.06 \pm 0.06 \pm 0.04$	LEES	11i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^0 \rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.12 \pm 0.17 \pm 0.02$	AUBERT	07z	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^{*+} \pi^+ \pi^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.07 \pm 0.07 \pm 0.04$	AUBERT,B	06u	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow \rho^0 K^*(892)^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.31 \pm 0.13 \pm 0.03$	DEL-AMO-SA...11D	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.20^{+0.32}_{-0.29} \pm 0.04$	AUBERT	03v	BABR Repl. by DEL-AMO-SANCHEZ 11D

 $A_{CP}(B^+ \rightarrow K^*(892)^+ f_0(980))$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.15 \pm 0.12 \pm 0.03$	DEL-AMO-SA...11D	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.34 \pm 0.21 \pm 0.03$	AUBERT,B	06G	BABR Repl. by DEL-AMO-SANCHEZ 11D

 $A_{CP}(B^+ \rightarrow a_1^+ K^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$+0.12 \pm 0.11 \pm 0.02$	AUBERT	08F	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow b_1^+ K^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.03 \pm 0.15 \pm 0.02$	AUBERT	08AG	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^*(892)^0 \rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.01 \pm 0.16 \pm 0.02$	AUBERT,B	06G	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow b_1^0 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.46 \pm 0.20 \pm 0.02$	AUBERT	07Bi	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^0 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.04 \pm 0.14$ OUR AVERAGE			
$0.014 \pm 0.168 \pm 0.002$	DUH	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.10 \pm 0.26 \pm 0.03$	1 AUBERT,BE	06c	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.13^{+0.23}_{-0.24} \pm 0.02$	LIN	07	BELL Repl. by DUH 13
$0.15 \pm 0.33 \pm 0.03$	2 AUBERT,BE	05E	BABR Repl. by AUBERT,BE 06c
1 Corresponds to 90% confidence range $-0.31 < A_{CP} < 0.54$ .			
2 Corresponds to 90% confidence range $-0.43 < A_{CP} < 0.68$ .			

 $A_{CP}(B^+ \rightarrow K_S^0 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.21 \pm 0.14 \pm 0.01$	AAIJ	13Bs	LHCB $pp$ at 7 TeV

 $A_{CP}(B^+ \rightarrow K^+ K_S^0 K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.04^{+0.04}_{-0.05} \pm 0.02$	LEES	12o	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.04 \pm 0.11 \pm 0.02$	1 AUBERT,B	04V	BABR Repl. by LEES 12o
1 Corresponds to 90% confidence range $-0.23 < A_{CP} < 0.15$ .			

# Meson Particle Listings

$B^\pm$

## $A_{CP}(B^+ \rightarrow K^+ K^- \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.118 \pm 0.022</math> OUR AVERAGE</b>			
$-0.123 \pm 0.017 \pm 0.014$	<sup>1</sup> AAIJ	14B0 LHCb	$pp$ at 7, 8 TeV
$0.00 \pm 0.10 \pm 0.03$	AUBERT	07Bb BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.141 \pm 0.040 \pm 0.019$	<sup>2</sup> AAIJ	14 LHCb	Repl. by AAIJ 14B0
<sup>1</sup> AAIJ 14B0 reports also $CP$ asymmetries in restricted regions of phase space.			
<sup>2</sup> AAIJ 14 reports $A_{CP}(B^+ \rightarrow K^+ K^- \pi^+) = -0.648 \pm 0.070 \pm 0.013 \pm 0.007$ in the Dalitz plot region of $m_{K^+ K^-}^2 < 1.5 \text{ GeV}^2/c^4$ . The third uncertainty is due to the $CP$ asymmetry of the $B^\pm \rightarrow J/\psi K^\pm$ reference mode uncertainty.			

## $A_{CP}(B^+ \rightarrow K^+ K^- K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.033 \pm 0.008</math> OUR AVERAGE</b>			
$-0.036 \pm 0.004 \pm 0.007$	<sup>1</sup> AAIJ	14B0 LHCb	$pp$ at 7, 8 TeV
$-0.017 \pm 0.019 \pm 0.014$	<sup>2</sup> LEES	120 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.043 \pm 0.009 \pm 0.008$	AAIJ	13AZ LHCb	Repl. by AAIJ 14B0
$-0.017 \pm 0.026 \pm 0.015$	AUBERT	060 BABR	Repl. by LEES 120
$0.02 \pm 0.07 \pm 0.03$	AUBERT	03M BABR	Repl. by AUBERT 060
<sup>1</sup> AAIJ 14B0 reports also $CP$ asymmetries in restricted regions of phase space.			
<sup>2</sup> All intermediate charmonium and charm resonances are removed, except of $\chi_{c0}$ .			

## $A_{CP}(B^+ \rightarrow \phi K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.024 \pm 0.028</math> OUR AVERAGE</b>	Error includes scale factor of 2.3.		
$0.017 \pm 0.011 \pm 0.006$	<sup>1</sup> AAIJ	150 LHCb	$pp$ at 7, 8 TeV
$0.128 \pm 0.044 \pm 0.013$	LEES	120 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.07 \pm 0.17 \pm 0.03$	ACOSTA	05J CDF	$p\bar{p}$ at 1.96 TeV
$0.01 \pm 0.12 \pm 0.05$	<sup>2</sup> CHEN	03B BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.022 \pm 0.021 \pm 0.009$	AAIJ	14A LHCb	Repl. by AAIJ 150
$0.00 \pm 0.08 \pm 0.02$	AUBERT	060 BABR	Repl. by LEES 120
$0.04 \pm 0.09 \pm 0.01$	<sup>3</sup> AUBERT	04A BABR	Repl. by AUBERT 060
$-0.05 \pm 0.20 \pm 0.03$	<sup>4</sup> AUBERT	02E BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained using $A_{CP}(B^\pm \rightarrow J/\psi K^\pm) = (0.3 \pm 0.6) \times 10^{-2}$ .			
<sup>2</sup> Corresponds to 90% confidence range $-0.20 < A_{CP} < 0.22$ .			
<sup>3</sup> Corresponds to 90% confidence range $-0.10 < A_{CP} < 0.18$ .			
<sup>4</sup> Corresponds to 90% confidence range $-0.37 < A_{CP} < 0.28$ .			

## $A_{CP}(B^+ \rightarrow X_0(1550) K^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.04 \pm 0.07 \pm 0.02</math></b>	<sup>1</sup> AUBERT	060 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Measured in the $B^+ \rightarrow K^+ K^- K^+$ decay.			

## $A_{CP}(B^+ \rightarrow K^{*+} K^+ K^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.11 \pm 0.08 \pm 0.03</math></b>	AUBERT,B	06U BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \phi K^*(892)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.01 \pm 0.08</math> OUR AVERAGE</b>			
$0.00 \pm 0.09 \pm 0.04$	AUBERT	07BA BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.02 \pm 0.14 \pm 0.03$	<sup>1</sup> CHEN	05A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.16 \pm 0.17 \pm 0.03$	AUBERT	03v BABR	Repl. by AUBERT 07BA
$-0.13 \pm 0.29 \pm 0.08$	<sup>2</sup> CHEN	03B BELL	Repl. by CHEN 05A
$-0.43 \pm 0.36 \pm 0.06$	<sup>3</sup> AUBERT	02E BABR	Repl. by AUBERT 03v
<sup>1</sup> Corresponds to 90% confidence range $-0.25 < A_{CP} < 0.22$ .			
<sup>2</sup> Corresponds to 90% confidence range $-0.64 < A_{CP} < 0.36$ .			
<sup>3</sup> Corresponds to 90% confidence range $-0.88 < A_{CP} < 0.18$ .			

## $A_{CP}(B^+ \rightarrow \phi(K\pi)_0^{*+})$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.04 \pm 0.15 \pm 0.04</math></b>	AUBERT	08B1 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \phi K_1(1270)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.15 \pm 0.19 \pm 0.05</math></b>	AUBERT	08B1 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \phi K_2^*(1430)^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.23 \pm 0.19 \pm 0.06</math></b>	AUBERT	08B1 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow K^+ \phi \phi)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.10 \pm 0.08 \pm 0.02</math></b>	<sup>1</sup> LEES	11A BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> $m_{\phi\phi} < 2.85 \text{ GeV}/c^2$ .			

## $A_{CP}(B^+ \rightarrow K^+[\phi\phi]_{\eta_c})$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.09 \pm 0.10 \pm 0.02</math></b>	<sup>1</sup> LEES	11A BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> $m_{\phi\phi}$ is consistent with $\eta_c$ mass [2.94, 3.02] GeV/ $c^2$ .			

## $A_{CP}(B^+ \rightarrow K^*(892)^+ \gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>+0.018 \pm 0.028 \pm 0.007</math></b>	AUBERT	09A0 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \eta K^+ \gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.12 \pm 0.07</math> OUR AVERAGE</b>			
$-0.09 \pm 0.10 \pm 0.01$	<sup>1</sup> AUBERT	09 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.16 \pm 0.09 \pm 0.06$	<sup>2</sup> NISHIDA	05 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.09 \pm 0.12 \pm 0.01$	<sup>1</sup> AUBERT,B	06M BABR	Repl. by AUBERT 09
<sup>1</sup> $m_{\eta K} < 3.25 \text{ GeV}/c^2$ .			
<sup>2</sup> $m_{\eta K} < 2.4 \text{ GeV}/c^2$			

## $A_{CP}(B^+ \rightarrow \phi K^+ \gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.13 \pm 0.11</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.		
$-0.03 \pm 0.11 \pm 0.08$	SAHOO	11A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.26 \pm 0.14 \pm 0.05$	AUBERT	07Q BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \rho^+ \gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.11 \pm 0.32 \pm 0.09</math></b>	TANIGUCHI	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

## $A_{CP}(B^+ \rightarrow \pi^+ \pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.03 \pm 0.04</math> OUR AVERAGE</b>			
$0.025 \pm 0.043 \pm 0.007$	DUH	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.03 \pm 0.08 \pm 0.01$	AUBERT	07Bc BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.07 \pm 0.06 \pm 0.01$	LIN	08 BELL	Repl. by DUH 13
$-0.01 \pm 0.10 \pm 0.02$	<sup>1</sup> AUBERT	05L BABR	Repl. by AUBERT 07Bc
$0.00 \pm 0.10 \pm 0.02$	<sup>2</sup> CHAO	05A BELL	Repl. by CHAO 04B
$-0.02 \pm 0.10 \pm 0.01$	<sup>3</sup> CHAO	04B BELL	Repl. by LIN 08
$-0.03 \pm 0.18 \pm 0.02$	<sup>4</sup> AUBERT	03L BABR	Repl. by AUBERT 05L
$0.30 \pm 0.30 \pm 0.06$	<sup>5</sup> CASEY	02 BELL	Repl. by CHAO 04B
<sup>1</sup> Corresponds to a 90% CL interval of $-0.19 < A_{CP} < 0.21$ .			
<sup>2</sup> Corresponds to a 90% CL interval of $-0.17 < A_{CP} < 0.16$ .			
<sup>3</sup> This corresponds to 90% CL interval of $-0.18 < A_{CP} < 0.14$ .			
<sup>4</sup> Corresponds to 90% confidence range $-0.32 < A_{CP} < 0.27$ .			
<sup>5</sup> Corresponds to 90% confidence range $-0.23 < A_{CP} < +0.86$ .			

## $A_{CP}(B^+ \rightarrow \pi^+ \pi^- \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.057 \pm 0.013</math> OUR AVERAGE</b>			
$0.058 \pm 0.008 \pm 0.011$	<sup>1</sup> AAIJ	14B0 LHCb	$pp$ at 7, 8 TeV
$0.032 \pm 0.044 \pm 0.040$	AUBERT	09L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.037 \pm 0.037$			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.117 \pm 0.021 \pm 0.011$	<sup>2</sup> AAIJ	14 LHCb	Repl. by AAIJ 14B0
$-0.007 \pm 0.077 \pm 0.025$	AUBERT,B	05G BABR	Repl. by AUBERT 09L
$-0.39 \pm 0.33 \pm 0.12$	AUBERT	03M BABR	Repl. by AUBERT 05G

<sup>1</sup> AAIJ 14B0 reports also  $CP$  asymmetries in restricted regions of phase space.  
<sup>2</sup> AAIJ 14 reports  $A_{CP}(B^+ \rightarrow \pi^+ \pi^- \pi^+) = 0.584 \pm 0.082 \pm 0.027 \pm 0.007$  in the Dalitz plot region of  $m_{\pi^+ \pi^-}^2 > 15 \text{ GeV}^2/c^4$  or  $m_{\pi^+ \pi^-}^2 < 0.4 \text{ GeV}^2/c^4$ . The third uncertainty is due to the  $CP$  asymmetry of the  $B^\pm \rightarrow J/\psi K^\pm$  reference mode uncertainty.

## $A_{CP}(B^+ \rightarrow \rho^0 \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.18 \pm 0.07 \pm 0.05</math></b>	AUBERT	09L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.074 \pm 0.120 \pm 0.035$	AUBERT,B	05G BABR	Repl. by AUBERT 09L
$-0.19 \pm 0.11 \pm 0.02$	AUBERT	04Z BABR	Repl. by AUBERT,B 05G

## $A_{CP}(B^+ \rightarrow f_2(1270) \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.41 \pm 0.25 \pm 0.18</math></b>	AUBERT	09L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.004 \pm 0.247 \pm 0.028$	AUBERT,B	05G BABR	Repl. by AUBERT 09L
$-0.032 \pm 0.032$			

## $A_{CP}(B^+ \rightarrow \rho^0(1450) \pi^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.06 \pm 0.28 \pm 0.23</math></b>	AUBERT	09L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.40 \pm 0.40$			



See key on page 601

## Meson Particle Listings

 $B^\pm$  $A_{CP}(B^+ \rightarrow f_0(1370)\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.72 \pm 0.15 \pm 0.16$	AUBERT	09L BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow \pi^+\pi^-\pi^+ \text{ nonresonant})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.14 \pm 0.14^{+0.18}_{-0.08}$	AUBERT	09L BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow \rho^+\pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.02 \pm 0.11</math> OUR AVERAGE</b>			
$-0.01 \pm 0.13 \pm 0.02$	AUBERT	07x BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.06 \pm 0.17^{+0.04}_{-0.05}$	ZHANG	05A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.24 \pm 0.16 \pm 0.06$	AUBERT	04z BABR	Repl. by AUBERT 07x

 $A_{CP}(B^+ \rightarrow \rho^+\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.05 \pm 0.05</math> OUR AVERAGE</b>			
$-0.054 \pm 0.055 \pm 0.010$	AUBERT	09G BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.00 \pm 0.22 \pm 0.03$	ZHANG	03B BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.12 \pm 0.13 \pm 0.10$	AUBERT,BE	06G BABR	Repl. by AUBERT 09G
$-0.19 \pm 0.23 \pm 0.03$	AUBERT	03v BABR	Repl. by AUBERT,BE 06G

 $A_{CP}(B^+ \rightarrow \omega\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.04 \pm 0.06</math> OUR AVERAGE</b>			
$-0.02 \pm 0.08 \pm 0.01$	AUBERT	07AE BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.02 \pm 0.09 \pm 0.01$	JEN	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.34 \pm 0.25$	<sup>1</sup> CHEN	00 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.01 \pm 0.10 \pm 0.01$	AUBERT,B	06E BABR	Repl. by AUBERT 07AE
$0.03 \pm 0.16 \pm 0.01$	AUBERT	04H BABR	Repl. by AUBERT,B 06E
$0.50 \pm 0.23^{+0.20}_{-0.20}$	<sup>2</sup> WANG	04A BELL	Repl. by JEN 06
$-0.01 \pm 0.29 \pm 0.03$	<sup>3</sup> AUBERT	02E BABR	Repl. by AUBERT 04H

<sup>1</sup> Corresponds to 90% confidence range  $-0.75 < A_{CP} < 0.07$ .<sup>2</sup> Corresponds to 90% CL interval  $-0.25 < A_{CP} < 0.41$ <sup>3</sup> Corresponds to 90% confidence range  $-0.50 < A_{CP} < 0.46$ . $A_{CP}(B^+ \rightarrow \omega\rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.20 \pm 0.09 \pm 0.02</math></b>	AUBERT	09H BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.04 \pm 0.18 \pm 0.02$	AUBERT,B	06T BABR	Repl. by AUBERT 09H
$0.05 \pm 0.26 \pm 0.02$	AUBERT	05o BABR	Repl. by AUBERT,B 06T

 $A_{CP}(B^+ \rightarrow \eta\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.14 \pm 0.07</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.		
$-0.19 \pm 0.06 \pm 0.01$	HOI	12 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.03 \pm 0.09 \pm 0.03$	AUBERT	09AV BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.08 \pm 0.10 \pm 0.01$	AUBERT	07AE BABR	Repl. by AUBERT 09AV
$-0.23 \pm 0.09 \pm 0.02$	CHANG	07B BELL	Repl. by HOI 12
$-0.13 \pm 0.12 \pm 0.01$	AUBERT,B	05K BABR	Repl. by AUBERT 07AE
$0.07 \pm 0.15 \pm 0.03$	CHANG	05A BELL	Repl. by CHANG 07B
$-0.44 \pm 0.18 \pm 0.01$	AUBERT	04H BABR	Repl. by AUBERT,B 05K

 $A_{CP}(B^+ \rightarrow \eta\rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.11 \pm 0.11</math> OUR AVERAGE</b>			
$0.13 \pm 0.11 \pm 0.02$	AUBERT	08AH BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.04 \pm 0.34^{+0.37}_{-0.32} \pm 0.01$	WANG	07B BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.02 \pm 0.18 \pm 0.02$	AUBERT,B	05K BABR	Repl. by AUBERT 08AH

 $A_{CP}(B^+ \rightarrow \eta'\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.16</math> OUR AVERAGE</b>			
$0.03 \pm 0.17 \pm 0.02$	AUBERT	09AV BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.20^{+0.37}_{-0.36} \pm 0.04$	SCHUEMANN	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.21 \pm 0.17 \pm 0.01$	AUBERT	07AE BABR	Repl. by AUBERT 09AV
$0.14 \pm 0.16 \pm 0.01$	AUBERT,B	05K BABR	Repl. by AUBERT 07AE

 $A_{CP}(B^+ \rightarrow \eta'\rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.26 \pm 0.17 \pm 0.02</math></b>	DEL-AMO-SA...10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.04 \pm 0.28 \pm 0.02$	<sup>1</sup> AUBERT	07E BABR	Repl. by DEL-AMO-SANCHEZ 10A

<sup>1</sup> Reports  $A_{CP}$  with the opposite sign convention. $A_{CP}(B^+ \rightarrow b_1^0\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>+0.05 \pm 0.16 \pm 0.02</math></b>	AUBERT	07Bi BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow \rho\bar{p}\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.04</math> OUR AVERAGE</b>			
$-0.02 \pm 0.05 \pm 0.02$	<sup>1</sup> WEI	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$+0.04 \pm 0.07 \pm 0.04$	AUBERT	07Av BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.16 \pm 0.22 \pm 0.01$	WANG	04 BELL	Repl. by WEI 08

<sup>1</sup> Requires  $m_{p\bar{p}} < 2.85 \text{ GeV}/c^2$ . $A_{CP}(B^+ \rightarrow \rho\bar{p}K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.04</math> OUR AVERAGE</b>	Error includes scale factor of 2.2.		
$0.021 \pm 0.020 \pm 0.004$	<sup>1</sup> AAIJ	14AF LHCb	$pp$ at 7, 8 TeV
$-0.17 \pm 0.10 \pm 0.02$	<sup>1</sup> WEI	08 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.16^{+0.07}_{-0.08} \pm 0.04$	<sup>1</sup> AUBERT,B	05L BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.047 \pm 0.036 \pm 0.007$	<sup>1</sup> AAIJ	13Au LHCb	Repl. by AAIJ 14AF
$-0.05 \pm 0.11 \pm 0.01$	WANG	04 BELL	Repl. by WEI 08

<sup>1</sup> Requires  $m_{p\bar{p}} < 2.85 \text{ GeV}/c^2$ . $A_{CP}(B^+ \rightarrow \rho\bar{p}K^*(892)^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.21 \pm 0.16</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.		
$-0.01 \pm 0.19 \pm 0.02$	CHEN	08c BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$+0.32 \pm 0.13 \pm 0.05$	AUBERT	07Av BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow \rho\bar{\Lambda}\gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>+0.17 \pm 0.16 \pm 0.05</math></b>	WANG	07c BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow \rho\bar{\Lambda}\pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>+0.01 \pm 0.17 \pm 0.04</math></b>	WANG	07c BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^+\ell^+\ell^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.02 \pm 0.08</math> OUR AVERAGE</b>			
$-0.03 \pm 0.14 \pm 0.01$	<sup>1</sup> LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.18 \pm 0.18 \pm 0.01$	AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$+0.04 \pm 0.10 \pm 0.02$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.07 \pm 0.22 \pm 0.02$	AUBERT,B	06J BABR	Repl. by AUBERT 09T
<sup>1</sup> Measured in the union of $0.10 < q^2 < 8.12 \text{ GeV}^2/c^4$ and $q^2 > 10.11 \text{ GeV}^2/c^4$ . LEES 12s reports also individual measurements $A_{CP}(B^+ \rightarrow K^+\ell^+\ell^-) = 0.02 \pm 0.18 \pm 0.01$ for $0.10 < q^2 < 8.12 \text{ GeV}^2/c^4$ and $A_{CP}(B^+ \rightarrow K^+\ell^+\ell^-) = -0.06^{+0.22}_{-0.21} \pm 0.01$ for $q^2 > 10.11 \text{ GeV}^2/c^4$ .			

 $A_{CP}(B^+ \rightarrow K^+e^+e^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>+0.14 \pm 0.14 \pm 0.03</math></b>	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^+ \rightarrow K^+\mu^+\mu^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.011 \pm 0.017</math> OUR AVERAGE</b>			
$0.012 \pm 0.017 \pm 0.001$	AAIJ	14AN LHCb	$pp$ at 7, 8 TeV
$-0.05 \pm 0.13 \pm 0.03$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.000 \pm 0.033 \pm 0.009$	AAIJ	13BN LHCb	Repl. by AAIJ 14AN

 $A_{CP}(B^+ \rightarrow \pi^+\mu^+\mu^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.11 \pm 0.12 \pm 0.01</math></b>	AAIJ	15AR LHCb	$pp$ at 7, 8 TeV

 $A_{CP}(B^+ \rightarrow K^{*+}\ell^+\ell^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.09 \pm 0.14</math> OUR AVERAGE</b>			
$0.01^{+0.26}_{-0.24} \pm 0.02$	AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.13^{+0.17}_{-0.16} \pm 0.01$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.03 \pm 0.23 \pm 0.03$	AUBERT,B	06J BABR	Repl. by AUBERT 09T

# Meson Particle Listings

$B^\pm$

$A_{CP}(B^+ \rightarrow K^* e^+ e^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.14^{+0.23}_{-0.22} \pm 0.02$	WEI	09A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^+ \rightarrow K^* \mu^+ \mu^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.12 \pm 0.24 \pm 0.02$	WEI	09A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

$\gamma(B^+ \rightarrow D^{(*)0} K^{(*)+})$   
For angle  $\gamma(\phi_3)$  of the CKM unitarity triangle, see the review on “CP Violation” in the Reviews section.

VALUE (°)	CL%	DOCUMENT ID	TECN	COMMENT
<b>70 ± 9 OUR AVERAGE</b>				
62 $^{+15}_{-14}$		<sup>1</sup> AAIJ	14BA	LHCB $pp$ at 7, 8 TeV
69 $^{+17}_{-16}$		<sup>2</sup> LEES	13B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
78.4 $^{+10.8}_{-11.6} \pm 9.6$		<sup>3</sup> POLUEKTOV	10	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
84 $^{+49}_{-42}$		<sup>4</sup> AAIJ	14BE	LHCB Repl. by AAIJ 14BA
72.6 $^{+9.7}_{-17.2}$		<sup>5</sup> AAIJ	13AK	LHCB $pp$ at 7 TeV
44 $^{+43}_{-38}$		<sup>6,7</sup> AAIJ	12AQ	LHCB Repl. by AAIJ 13AK
77.3 $^{+15.1}_{-14.9} \pm 5.9$		<sup>7,8</sup> AIHARA	12	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
68 $\pm 14 \pm 5$		<sup>9</sup> DEL-AMO-SA...10F	BABR	Repl. by LEES 13B
7 to 173	95	<sup>10</sup> DEL-AMO-SA...10G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
76 $^{+22}_{-23} \pm 7.1$		<sup>11</sup> AUBERT	08AL	BABR Repl. by DEL-AMO-SANCHEZ 10F
53 $^{+15}_{-18} \pm 10$		<sup>12</sup> POLUEKTOV	06	BELL Repl. by POLUEKTOV 10
70 $\pm 31 \pm 18$		<sup>13</sup> AUBERT,B	05Y	BABR Repl. by AUBERT 08AL
77 $^{+17}_{-19} \pm 17$		<sup>14</sup> POLUEKTOV	04	BELL Repl. by POLUEKTOV 06

- <sup>1</sup> Uses binned Dalitz plot analysis of  $B^+ \rightarrow DK^+$  decays, with  $D \rightarrow K_S^0 \pi^+ \pi^-$  and  $D \rightarrow K_S^0 K^+ K^-$ . Strong phase measurements from CLEO-c (LIBBY 10) of the  $D$  decay over the Dalitz plot are used as input. Solution that satisfies  $0 < \gamma < 180$  is chosen.
- <sup>2</sup> Reports combination of published measurements using GGSZ, GLW, and ADS methods. Reports also  $2\sigma$  range of  $41\text{--}102^\circ$  and a  $5.9\sigma$  significance for  $\gamma(B^+ \rightarrow D^{(*)0} K^{(*)+}) \neq 0$  hypothesis.
- <sup>3</sup> Uses Dalitz plot analysis of  $\overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow D^{(*)} K^+$  modes. The corresponding two standard deviation interval for  $\gamma$  is  $54.2^\circ < \gamma < 100.5^\circ$ . CP conservation in the combined result is ruled out with a significance of 3.5 standard deviations.
- <sup>4</sup> AAIJ 14BE uses model-dependent analysis of  $D \rightarrow K_S^0 \pi^+ \pi^-$  amplitudes. The model is the same as in DEL-AMO-SANCHEZ 10F.
- <sup>5</sup> Presents a confidence region  $55.4^\circ < \gamma < 82.3^\circ$  at 68% CL with best fit value  $72.6^\circ$  and includes both statistical and systematic uncertainties. The corresponding 95% CL is  $40.2^\circ < \gamma < 92.7^\circ$ . The value is determined from combination of measurements using  $D$  meson decaying to  $K^+ K^-$ ,  $\pi^+ \pi^-$ ,  $K^\pm \pi^\mp$ ,  $K_S^0 \pi^+ \pi^-$ ,  $K_S^0 K^+ K^-$ , and  $K^\pm \pi^\mp \pi^\pm \pi^\mp$ . Combines  $B^\pm \rightarrow DK^\pm$  and  $B^\pm \rightarrow D\pi^\pm$ .
- <sup>6</sup> Reports combined statistical and systematic uncertainties.
- <sup>7</sup> Uses binned Dalitz plot of  $\overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays from  $B^+ \rightarrow \overline{D}^0 K^+$ . Measurement of strong phases in  $\overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  Dalitz plot from LIBBY 10 is used as input.
- <sup>8</sup> We combined the systematics in quadrature. The authors report separately the contribution to the systematic uncertainty due to the uncertainty on the bin-averaged strong phase difference between  $D^0$  and  $\overline{D}^0$  amplitudes.
- <sup>9</sup> Uses Dalitz plot analysis of  $\overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$ ,  $K_S^0 K^+ K^-$  decays from  $B^+ \rightarrow D^{(*)} K^+$ ,  $DK^{*+}$  modes. The corresponding two standard deviation interval for  $\gamma$  is  $39^\circ < \gamma < 98^\circ$ . CP conservation in the combined result is ruled out with a significance of 3.5 standard deviations.
- <sup>10</sup> Reports confidence intervals for the CKM angle  $\gamma$  from the measured values of the GLW parameters using  $B^\pm \rightarrow DK^\pm$  decays from  $D$  mesons decaying to non-CP( $K\pi$ ), CP-even ( $K^+ K^-$ ,  $\pi^+ \pi^-$ ), and CP-odd ( $K_S^0 \pi^0$ ,  $K_S^0 \omega$ ) states.
- <sup>11</sup> Uses Dalitz plot analysis of  $\overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  and  $\overline{D}^0 \rightarrow K_S^0 K^+ K^-$  decays coming from  $B^\pm \rightarrow D^{(*)} K^{(*)\pm}$  modes. The corresponding two standard deviation interval is  $29^\circ < \gamma < 122^\circ$ .
- <sup>12</sup> Uses a Dalitz plot analysis of the  $\overline{D}^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays; Combines the  $DK^+$ ,  $D^* K^+$  and  $DK^{*+}$  modes. The corresponding two standard deviations interval for gamma is  $8^\circ < \gamma < 111^\circ$ .
- <sup>13</sup> Uses a Dalitz plot analysis of neutral  $D \rightarrow K_S^0 \pi^+ \pi^-$  decays coming from  $B^\pm \rightarrow DK^\pm$  and  $B^\pm \rightarrow D^{*0} K^\pm$  followed by  $D^{*0} \rightarrow D\pi^0$ ,  $D\gamma$ . The corresponding two standard deviations interval for gamma is  $12^\circ < \gamma < 137^\circ$ . AUBERT,B 05Y also reports the amplitude ratios and the strong phases.
- <sup>14</sup> Uses a Dalitz plot analysis of the 3-body  $D \rightarrow K_S^0 \pi^+ \pi^-$  decays coming from  $B^\pm \rightarrow DK^\pm$  and  $B^\pm \rightarrow D^* K^\pm$  followed by  $D^* \rightarrow D\pi^0$ ; here we use  $D$  to denote that the neutral  $D$  meson produced in the decay is an admixture of  $D^0$  and  $\overline{D}^0$ . The corresponding two standard deviations interval for  $\gamma$  is  $26^\circ < \gamma < 126^\circ$ . POLUEKTOV 04 also reports the amplitude ratios and the strong phases.

$\gamma(B^+ \rightarrow DK^+ \pi^- \pi^+, D\pi^+ \pi^- \pi^+)$			
VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>74 <math>^{+20}_{-19}</math></b>	AAIJ	15Bc	LHCB $pp$ at 7, 8 TeV

## PARTIAL BRANCHING FRACTIONS

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (q^2 < 2.0 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.4 ± 0.5 OUR AVERAGE</b>			
1.37 $^{+0.60}_{-0.58}$	AAIJ	12AH	LHCB $pp$ at 7 TeV
1.30 $\pm 0.98 \pm 0.14$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.1 ± 0.5 OUR AVERAGE</b>			
1.24 $^{+0.60}_{-0.55}$	AAIJ	12AH	LHCB $pp$ at 7 TeV
0.71 $\pm 1.00 \pm 0.15$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (4.3 < q^2 < 8.68 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.4 <math>^{+0.8}_{-0.7}</math> OUR AVERAGE</b>			
2.50 $^{+0.88}_{-0.74}$	AAIJ	12AH	LHCB $pp$ at 7 TeV
1.71 $\pm 1.58 \pm 0.49$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.1 ± 0.6 OUR AVERAGE</b>			
2.13 $^{+0.72}_{-0.66}$	AAIJ	12AH	LHCB $pp$ at 7 TeV
1.97 $\pm 0.99 \pm 0.22$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (14.18 < q^2 < 16.0 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.86 <math>^{+0.40}_{-0.32}</math> OUR AVERAGE</b>			
1.00 $^{+0.47}_{-0.38}$	AAIJ	12AH	LHCB $pp$ at 7 TeV
0.52 $\pm 0.61 \pm 0.09$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (15.0 < q^2 < 19.0 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.58 <math>^{+0.32}_{-0.29} \pm 0.11</math></b>	<sup>1</sup> AAIJ	14M	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Uses $B(B^+ \rightarrow J/\psi(1S) K^*(892)^+) = (1.431 \pm 0.027 \pm 0.090) \times 10^{-3}$ for normalization and $\mu^+ \mu^-$ as a lepton pair.			

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (q^2 > 16.0 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.3 ± 0.4 OUR AVERAGE</b>			
1.25 $\pm 0.46$	AAIJ	12AH	LHCB $pp$ at 7 TeV
1.57 $\pm 0.96 \pm 0.17$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.8 ± 0.4 OUR AVERAGE</b>			
1.79 $^{+0.41}_{-0.37} \pm 0.13$	<sup>1</sup> AAIJ	14M	LHCB $pp$ at 7, 8 TeV
2.57 $\pm 1.61 \pm 0.40$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.90 $^{+0.90}_{-0.85}$	AAIJ	12AH	LHCB Repl. by AAIJ 14M
<sup>1</sup> Uses $B(B^+ \rightarrow J/\psi(1S) K^*(892)^+) = (1.431 \pm 0.027 \pm 0.090) \times 10^{-3}$ for normalization and $\mu^+ \mu^-$ as a lepton pair. Measured in $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ .			

$B(B^+ \rightarrow K^{*+} \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.01 <math>\pm 1.39 \pm 0.27</math></b>	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^+ \ell^+ \ell^-) (q^2 < 2.0 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.51 ± 0.08 OUR AVERAGE</b>	Error includes scale factor of 1.5.		
0.556 $\pm 0.053 \pm 0.027$	<sup>1</sup> AAIJ	13H	LHCB $pp$ at 7 TeV
0.36 $\pm 0.11 \pm 0.03$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV
<sup>1</sup> Measured in $0.05 < q^2 < 2.0 \text{ GeV}^2/c^4$ range.			

$B(B^+ \rightarrow K^+ \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.60 ± 0.07 OUR AVERAGE</b>	Error includes scale factor of 1.3.		
0.573 $\pm 0.053 \pm 0.023$	AAIJ	13H	LHCB $pp$ at 7 TeV
0.80 $\pm 0.15 \pm 0.05$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^+ \ell^+ \ell^-) (4.3 < q^2 < 8.68 \text{ GeV}^2/c^4)$			
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.03 ± 0.07 OUR AVERAGE</b>			
1.003 $\pm 0.070 \pm 0.039$	AAIJ	13H	LHCB $pp$ at 7 TeV
1.18 $\pm 0.19 \pm 0.09$	AALTONEN	11Ai	CDF $p\overline{p}$ at 1.96 TeV

$B(B^+ \rightarrow K^+ \ell^+ \ell^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.58 \pm 0.05</math> OUR AVERAGE</b>			
$0.565 \pm 0.050 \pm 0.022$	AAIJ	13H LHCb	$pp$ at 7 TeV
$0.68 \pm 0.12 \pm 0.05$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (14.18 < q^2 < 16.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.40 \pm 0.05</math> OUR AVERAGE</b>			Error includes scale factor of 1.4.
$0.377 \pm 0.036 \pm 0.015$	AAIJ	13H LHCb	$pp$ at 7 TeV
$0.53 \pm 0.10 \pm 0.03$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (16.0 < q^2 < 18.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.354 \pm 0.036 \pm 0.018</math></b>	AAIJ	13H LHCb	$pp$ at 7 TeV

 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (18.0 < q^2 < 22.0 \text{ GeV}^2/c^4)$ 

$F_H$  is a fractional contribution of (pseudo) scalar and tensor amplitudes to the decay width in the massless muon approximation.

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.312 \pm 0.040 \pm 0.016</math></b>	AAIJ	13H LHCb	$pp$ at 7 TeV

 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (15.0 < q^2 < 22.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.85 \pm 0.03 \pm 0.04</math></b>	<sup>1</sup> AAIJ	14M LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Uses  $B(B^+ \rightarrow J/\psi(1S) K^+) = (0.998 \pm 0.014 \pm 0.040) \times 10^{-3}$  for normalization and  $\mu^+ \mu^-$  as a lepton pair.

 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (16.0 < q^2 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.48 \pm 0.11 \pm 0.03</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

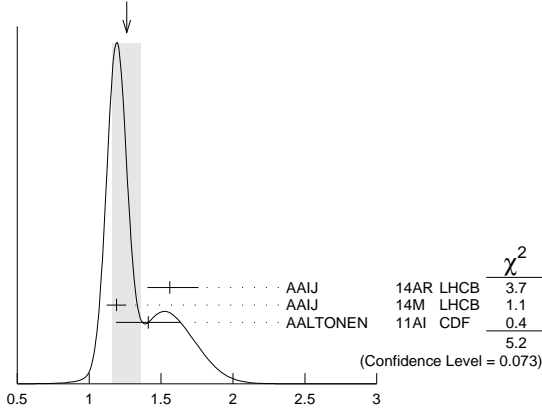
 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.26 \pm 0.10</math> OUR AVERAGE</b>			Error includes scale factor of 1.6. See the ideogram below.
$1.56^{+0.19}_{-0.15}^{+0.06}_{-0.04}$	<sup>1</sup> AAIJ	14AR LHCb	$pp$ at 7, 8 TeV
$1.19 \pm 0.034 \pm 0.059$	<sup>2</sup> AAIJ	14M LHCb	$pp$ at 7, 8 TeV
$1.41 \pm 0.20 \pm 0.10$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.205 \pm 0.085 \pm 0.070$	AAIJ	13H LHCb	Repl. by AAIJ 14M

<sup>1</sup> Measured by taking the ratio of the branching fraction from  $B^+ \rightarrow K^+ e^+ e^-$  and  $B^+ \rightarrow J/\psi(e^+ e^-) K^+$  decays and multiplying it by the measured value of  $B^+ \rightarrow J/\psi K^+$  and  $J/\psi \rightarrow e^+ e^-$  as in PDG 12 update. The branching fraction of  $B^+ \rightarrow K^+ e^+ e^-$  is determined in the region  $1 < q^2 < 6 \text{ GeV}^2/c^4$ .

<sup>2</sup> Uses  $B(B^+ \rightarrow J/\psi(1S) K^+) = (0.998 \pm 0.014 \pm 0.040) \times 10^{-3}$  for normalization and  $\mu^+ \mu^-$  for leptons. Measured for  $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ .

WEIGHTED AVERAGE  
1.26±0.10 (Error scaled by 1.6)



$B(B^+ \rightarrow K^+ \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4) \text{ (units } 10^{-7})$

 $B(B^+ \rightarrow K^+ \mu^+ \mu^-) / B(B^+ \rightarrow K^+ e^+ e^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.745^{+0.090}_{-0.074} \pm 0.036</math></b>	<sup>1</sup> AAIJ	14AR LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> The ratio is determined using the ratio of the relative branching fractions of the decays  $B^+ \rightarrow K^+ \ell^+ \ell^-$  and  $B^+ \rightarrow J/\psi(\rightarrow \ell^+ \ell^-) K^+$ , with  $\ell = e, \mu$ .

 $B(B^+ \rightarrow K^+ \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.13 \pm 0.19 \pm 0.08</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

 $B(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-) (1.00 < q^2 < 6.00 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.38^{+0.15}_{-0.14} \pm 0.08</math></b>	AAIJ	14AZ LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-) (0.10 < q^2 < 2.00 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.33^{+0.13}_{-0.12} \pm 0.09</math></b>	AAIJ	14AZ LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-) (2.00 < q^2 < 4.30 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.38^{+0.94}_{-0.87} \pm 0.35</math></b>	AAIJ	14AZ LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-) (4.30 < q^2 < 8.68 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.01^{+0.12}_{-0.13} \pm 0.09</math></b>	AAIJ	14AZ LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.07^{+0.94}_{-0.89} \pm 0.47</math></b>	AAIJ	14AZ LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow K^+ \pi^+ \pi^- \mu^+ \mu^-) (14.18 < q^2 < 19.00 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.48^{+0.39}_{-0.29} \pm 0.05</math></b>	AAIJ	14AZ LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow \pi^+ \mu^+ \mu^-) / B(B^+ \rightarrow K^+ \mu^+ \mu^-) (1.00 < q^2 < 6.00 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.8 \pm 0.9 \pm 0.1</math></b>	AAIJ	15AR LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow \pi^+ \mu^+ \mu^-) (1.00 < q^2 < 6.00 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-9}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>4.55^{+1.05}_{-1.00} \pm 0.15</math></b>	AAIJ	15AR LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow \pi^+ \mu^+ \mu^-) (15.00 < q^2 < 22.00 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-9}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.29^{+0.84}_{-0.70} \pm 0.07</math></b>	AAIJ	15AR LHCb	$pp$ at 7, 8 TeV

 $B(B^+ \rightarrow \pi^+ \mu^+ \mu^-) / B(B^+ \rightarrow K^+ \mu^+ \mu^-) (15.0 < q^2 < 22.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.7 \pm 0.8 \pm 0.1</math></b>	AAIJ	15AR LHCb	$pp$ at 7, 8 TeV

 $A_{FB}(B^+ \rightarrow K^+ \mu^+ \mu^-) (1.1 < q^2 < 6.0 \text{ GeV}^2/c^4)$ 

$A_{FB}$  is the forward-backward angular asymmetry of the lepton pair in  $B \rightarrow K^{(*)} \ell^+ \ell^-$  decay as defined in  $B^+, B^0$  admixture particle listings.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.005 \pm 0.015 \pm 0.010</math></b>	<sup>1</sup> AAIJ	14O LHCb	$pp$ at 7, 8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.02^{+0.05}_{-0.03}^{+0.02}_{-0.01}$	AAIJ	13H LHCb	Repl. by AAIJ 14O
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<sup>1</sup> AAIJ 14O reports 68% C.L. interval, which we encode as midpoint with uncertainty as half of the width of interval.

 $A_{FB}(B^+ \rightarrow K^+ \mu^+ \mu^-) (15.0 < q^2 < 22.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.015 \pm 0.015 \pm 0.01</math></b>	<sup>1</sup> AAIJ	14O LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> AAIJ 14O reports 68% C.L. interval, which we encode as midpoint with uncertainty as half of the width of interval.

 $F_H(B^+ \rightarrow K^+ \mu^+ \mu^-) (1.1 < q^2 < 6.0 \text{ GeV}^2/c^4)$ 

$F_H$  is a fractional contribution of (pseudo) scalar and tensor amplitudes to the decay width in the massless muon approximation.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.03 \pm 0.03 \pm 0.02</math></b>	<sup>1</sup> AAIJ	14O LHCb	$pp$ at 7, 8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.05^{+0.08}_{-0.05}^{+0.04}_{-0.02}$	AAIJ	13H LHCb	Repl. by AAIJ 14O
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<sup>1</sup> AAIJ 14O reports 68% C.L. interval, which we encode as midpoint with uncertainty as half of the width of interval.

 $F_H(B^+ \rightarrow K^+ \mu^+ \mu^-) (15.0 < q^2 < 22.0 \text{ GeV}^2/c^4)$ 

$F_H$  is a fractional contribution of (pseudo) scalar and tensor amplitudes to the decay width in the massless muon approximation.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.035 \pm 0.035 \pm 0.02</math></b>	<sup>1</sup> AAIJ	14O LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> AAIJ 14O reports 68% C.L. interval, which we encode as midpoint with uncertainty as half of the width of interval.

Meson Particle Listings

$B^\pm$

FORWARD-BACKWARD ASYMMETRIES

The forward-backward assymetry is defined as  $A_{FB} = [N(q_{FB} > 0) - N(q_{FB} < 0)] / [N(q_{FB} > 0) + N(q_{FB} < 0)]$ , where  $q_{FB} = -q_B \cdot \text{sgn}(\eta_B)$  with  $q_B$  as the  $B$  hadron electric charge,  $\eta_B$  as its pseudorapidity, and  $\text{sgn}(\eta_B)$  as a sign function of  $\eta_B$ .

$A_{FB}(B^\pm \rightarrow J/\psi K^\pm)$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
$-0.24 \pm 0.41 \pm 0.19$	ABAZOV	15	D0 $p\bar{p}$ at 1.96 TeV

$B^\pm$  REFERENCES

LEES	16	PRL 116 041801	J.P. Lees <i>et al.</i>	(BABAR Collab.)
AAU	15AR	JHEP 1510 034	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	15BC	PR D92 112005	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	15O	PRL 115 051801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	15V	PR D91 092002	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	15W	PR D91 112014	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	15	PRL 114 051803	V.M. Abazov <i>et al.</i>	(D0 Collab.)
BALA	15	PR D91 051101	A. Bala <i>et al.</i>	(BELLE Collab.)
CHOI	15A	PR D91 092011	S.-K. Choi <i>et al.</i>	(BELLE Collab.)
GOH	15	PR D91 071101	Y.M. Goh <i>et al.</i>	(BELLE Collab.)
HELLER	15	PR D91 112009	A. Heller <i>et al.</i>	(BELLE Collab.)
KRONENBIT...	15	PR D92 051102	B. Kronenbitter <i>et al.</i>	(BELLE Collab.)
LEES	15	PR D91 012003	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	15C	PR D91 052002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
VINOKUROVA	15	JHEP 1506 132	A. Vinokurova <i>et al.</i>	(BELLE Collab.)
WIECHCZYN...	15	PR D91 032008	J. Wiechczynski <i>et al.</i>	(BELLE Collab.)
YOOK	15	PR D91 052016	Y. Yook <i>et al.</i>	(BELLE Collab.)
AAU	14	PRL 112 011801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14A	PL B728 85	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14AC	PRL 112 131802	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14AF	PRL 113 141801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14AN	JHEP 1409 177	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14AR	PRL 113 151601	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14AZ	JHEP 1410 064	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14BA	JHEP 1410 097	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14BE	NP B868 169	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14BO	PR D90 112004	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14E	JHEP 1404 114	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14M	JHEP 1406 133	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14O	JHEP 1405 082	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	14V	PL B733 36	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	14A	PR D89 012004	V.M. Abazov <i>et al.</i>	(D0 Collab.)
CHOBANOVA	14	PR D90 012002	V. Chobanova <i>et al.</i>	(BELLE Collab.)
IWASHITA	14	PTEP 2014 043C01	T. Iwashita <i>et al.</i>	(BELLE Collab.)
LEES	14A	PR D89 011102	J.P. Lees <i>et al.</i>	(BABAR Collab.)
TJEN	14	PR D89 011101	K.-J. Tien <i>et al.</i>	(BELLE Collab.)
AAU	13AE	PL B723 44	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13AF	PL B726 151	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13AP	PR D87 092007	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13AU	PR D88 052015	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13AZ	PRL 111 101801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13BC	PRL 111 112003	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13BN	PRL 111 151801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13BS	PL B726 646	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13H	JHEP 1302 105	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13J	JHEP 1302 043	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13S	EPJ C73 2462	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	13Z	JHEP 1309 006	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	13M	PRL 110 241801	V.M. Abazov <i>et al.</i>	(D0 Collab.)
DUH	13	PR D87 031103	Y. T. Duh <i>et al.</i>	(BELLE Collab.)
HARA	13	PRL 110 131801	K. Hara <i>et al.</i>	(BELLE Collab.)
LEES	13A	PR D87 032004	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13B	PR D87 052015	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13I	PR D87 112005	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13K	PR D88 031102	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13M	PR D88 032012	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	13T	PR D88 072006	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LUTZ	13	PR D87 111103	O. Lutz <i>et al.</i>	(BELLE Collab.)
NAYAK	13	PR D88 091104	M. Nayak <i>et al.</i>	(BELLE Collab.)
SIBIDANOV	13	PR D88 032005	A. Sibidanov <i>et al.</i>	(LHCb Collab.)
AAU	12AA	PR D85 091102	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12AC	PR D85 091105	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12AD	PR D85 112004	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12AH	JHEP 1207 133	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12AQ	PL B718 43	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12AY	JHEP 1212 125	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12C	PRL 108 101601	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12E	PL B708 241	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12L	EPJ C72 2118	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12M	PL B712 203	R. Aaij <i>et al.</i>	(LHCb Collab.)
Also		PL B713 351 (errat.)	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	12T	PRL 108 161801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AIHARA	12	PR D85 112014	H. AiHara <i>et al.</i>	(BELLE Collab.)
DEL-AMO-SA...	12	PR D85 092017	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
HOI	12	PRL 108 031801	C.-T. Hoi <i>et al.</i>	(BELLE Collab.)
KIM	12A	PR D86 031101	J.H. Kim <i>et al.</i>	(BELLE Collab.)
LEES	12AA	PR D86 092004	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12B	PR D85 052003	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12D	PRL 109 101802	J.P. Lees <i>et al.</i>	(BABAR Collab.)
Also		PR D88 072012	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12J	PR D85 071103	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12O	PR D85 112010	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12P	PR D86 012004	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12S	PR D86 032012	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12Z	PR D86 091102	J.P. Lees <i>et al.</i>	(BABAR Collab.)
PDG	12	PR D86 010001	J. Beringer <i>et al.</i>	(PDG Collab.)
STYPULA	12	PR D86 072007	J. Stypula <i>et al.</i>	(BELLE Collab.)
AAU	11E	PR D84 092001	R. Aaij <i>et al.</i>	(LHCb Collab.)
Also		PR D85 039904 (errat.)	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	11	PRL 106 121804	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11AI	PRL 107 201802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11AJ	PR D84 091504	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11B	PR D83 032009	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11L	PRL 106 161801	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AUSHEV	11	PR D83 051102	T. Aushev <i>et al.</i>	(BELLE Collab.)
BHARDWAJ	11	PRL 107 091803	V. Bhardwaj <i>et al.</i>	(BELLE Collab.)
CHEN	11F	PR D84 071501	P. Chen <i>et al.</i>	(BELLE Collab.)
CHOI	11	PR D84 052004	S.-K. Choi <i>et al.</i>	(BELLE Collab.)
DEL-AMO-SA...	11B	PR D83 032004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	11C	PR D83 032007	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	11D	PR D83 051101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	11F	PR D83 052011	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	11K	PR D83 091101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)

DEL-AMO-SA...	11L	PRL 107 041804	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
GULER	11	PR D83 032005	H. Guler <i>et al.</i>	(BELLE Collab.)
HORII	11	PRL 106 231803	Y. Horii <i>et al.</i>	(BELLE Collab.)
LEES	11A	PR D84 012001	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	11D	PR D84 012002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	11I	PR D84 092007	J.P. Lees <i>et al.</i>	(BABAR Collab.)
SAHOO	11A	PR D84 071101	H. Sahoo <i>et al.</i>	(BELLE Collab.)
SEON	11	PR D84 071106	O. Seon <i>et al.</i>	(BELLE Collab.)
VINOKUROVA	11	PL B706 139	A. Vinokurova <i>et al.</i>	(BELLE Collab.)
AALTONEN	10A	PR D81 031105	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AUBERT	10	PRL 104 011802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	10D	PR D81 052009	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	10E	PR D81 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUSHEV	10	PR D81 031103	T. Aushev <i>et al.</i>	(BELLE Collab.)
BOZEK	10	PR D82 072005	A. BOZEK <i>et al.</i>	(BELLE Collab.)
DEL-AMO-SA...	10A	PR D82 011502	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10B	PR D82 011101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10F	PRL 105 121801	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10G	PR D82 027004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10H	PR D82 072006	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10I	PR D82 091101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10K	PR D82 092006	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...	10Q	PR D82 112002	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
HARA	10	PR D82 071101	K. Hara <i>et al.</i>	(BELLE Collab.)
LIBBY	10	PR D82 112006	J. Libby <i>et al.</i>	(CLEO Collab.)
POLUEKTOV	10	PR D81 120002	A. Poluektov <i>et al.</i>	(BELLE Collab.)
SAKAI	10	PR D82 091104	K. Sakai <i>et al.</i>	(BELLE Collab.)
WEDD	10	PR D81 111104	R. Wedd <i>et al.</i>	(BELLE Collab.)
AALTONEN	09B	PR D79 011104	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	09Y	PR D79 111102	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABULENCIA	09	PR D79 112003	A. Abulencia <i>et al.</i>	(CDF Collab.)
AUBERT	09	PR D79 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09A	PR D79 012002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AA	PR D79 112001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AB	PR D79 112004	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AF	PR D80 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AJ	PR D80 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AO	PRL 103 211802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AT	PR D80 111105	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09AV	PR D80 112002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09B	PRL 102 132001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09F	PR D79 051102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09G	PRL 102 141802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09H	PR D79 052005	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09J	PR D79 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09L	PR D79 072006	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09Q	PR D79 052011	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09S	PR D79 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09T	PRL 102 091803	B. Aubert <i>et al.</i>	(BABAR Collab.)
Also		EPAPS Document No. E-PR/TAO-102-060910		(BABAR Collab.)
AUBERT	09V	PR D79 091101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	09Y	PRL 103 051803	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHANG	09	PR D79 052006	Y.-W. Chang <i>et al.</i>	(BELLE Collab.)
CHEN	09C	PR D80 111103	P. Chen <i>et al.</i>	(BELLE Collab.)
LIU	09	PR D79 071102	C. Liu <i>et al.</i>	(BELLE Collab.)
WEI	09A	PRL 103 171801	J.-T. Wei <i>et al.</i>	(BELLE Collab.)
Also		EPAPS Supplement EPAPS_appendix.pdf		(BELLE Collab.)
WIECHCZYN...	09	PR D80 052005	J. Wiechczynski <i>et al.</i>	(BELLE Collab.)
ABAZOV	08O	PRL 100 211802	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ADACHI	08	PR D77 091101	I. Adachi <i>et al.</i>	(BELLE Collab.)
AUBERT	08A	PR D77 011101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AA	PR D77 111102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AB	PR D78 012006	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AD	PR D77 091104	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AG	PR D78 011104	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AH	PR D78 011107	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AI	PR D78 012004	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AL	PR D78 034023	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AT	PRL 100 231803	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08AV	PRL 101 081801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08B	PR D77 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BC	PR D78 072007	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BD	PR D78 091101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BE	PR D78 091102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BF	PR D78 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BH	PR D78 112001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BI	PRL 101 161801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BK	PRL 101 201801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BL	PRL 101 261802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BN	PR D78 112003	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08D	PR D77 011107	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08F	PRL 100 051803	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08G	PRL 100 171803	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08H	PR D77 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08N	PRL 100 021801	B. Aubert <i>et al.</i>	(BABAR Collab.)
Also		PR D79 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08Q	PRL 100 151802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08W	PRL 101 082001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08X	PRL 101 091801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08Y	PR D77 111101	B. Aubert <i>et al.</i>	(BABAR Collab.)
BHARDWAJ	08	PR D78 051104	V. Bhardwaj <i>et al.</i>	(BELLE Collab.)
BRODZICKA	08	PRL 100 092001	J. Brodzicka <i>et al.</i>	(BELLE Collab.)
CHEN	08C	PRL 100 251801	J.-H. Chen <i>et al.</i>	(BELLE Collab.)
HORII	08	PR D78 071901	Y. Horii <i>et al.</i>	(BELLE Collab.)
IWABUCHI	08	PRL 101 041601	M. Iwabuchi <i>et al.</i>	(BELLE Collab.)
LIN	08	NAT 452 3153	S.-W. Lin <i>et al.</i>	(BELLE Collab.)
LIVENTSEV	08	PR D77 091503	D. Liventsev <i>et al.</i>	(BELLE Collab.)
PL	08	PL B667 80	A. Pl <i>et al.</i>	(CDF Collab.)
TANIGUCHI	08	PRL 101 111801	N. Taniguchi <i>et al.</i>	(BELLE Collab.)
WEI	08	PL B659 80	J.-T. Wei <i>et al.</i>	(BELLE Collab.)
WEI	08A	PR D78 011101	J.-T. Wei <i>et al.</i>	(BELLE Collab.)
WICHT	08	PL B662 323	J. Wicht <i>et al.</i>	(BELLE Collab.)
ADAM	07	PRL 99 041802	N.E. Adam <i>et al.</i>	(CLEO Collab.)
Also		PR D76 012007	D.M. Asner <i>et al.</i>	(CLEO Collab.)
AUBERT	07AC	PR D76 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AE	PR D76 031103	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AF	PR 99 051801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AN	PR D76 051002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AT	PR D76 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AR	PR D76 071103	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AV	PR D76 092004	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07AZ	PRL 99 201801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BA	PRL 99 201802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BB	PRL 99 221801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BC	PR D76 091102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BI	PRL 99 241803	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BJ	PRL 99 251801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BK	PRL 99 261801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07BN	PR D76 111101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07H	PR D8 051802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07H	PR D75 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07L	PRL 98 151802	B. Aubert <i>et al.</i>	(BABAR Collab.)

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## Meson Particle Listings

 $B^{\pm}$ 

AUBERT	07M	PRL 98 171801	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	04Q	PR D69 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07N	PR D75 072002	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	04T	PR D69 071103	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07Q	PR D75 051102	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	04Y	PRL 93 041801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07R	PRL 98 211804	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	04Z	PRL 93 051802	B. Aubert <i>et al.</i>	(BABAR Collab.)
Also		PRL 100 189903E	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT,B	04B	PR D70 011101	B. Aubert <i>et al.</i>	(BABAR Collab.)
Also		PRL 100 199905E	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT,B	04D	PR D70 032006	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07X	PR D75 091103	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT,B	04L	PRL 93 131804	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	07Z	PR D76 011103	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT,B	04P	PR D70 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHANG	07B	PR D75 071104	P. Chang <i>et al.</i>	(BELLE Collab.)	AUBERT,B	04S	PRL 93 181801	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHEN	07D	PRL 99 221802	K.-F. Chen <i>et al.</i>	(BELLE Collab.)	AUBERT,B	04U	PR D70 091105	B. Aubert <i>et al.</i>	(BABAR Collab.)
HOKUUE	07	PL B648 139	T. Hokuue <i>et al.</i>	(BELLE Collab.)	AUBERT,B	04V	PRL 93 181805	B. Aubert <i>et al.</i>	(BABAR Collab.)
LIN	07	PRL 98 181804	S.-W. Lin <i>et al.</i>	(BELLE Collab.)	AUBERT,BE	04	PR D70 111102	B. Aubert <i>et al.</i>	(BABAR Collab.)
LIN	07A	PRL 99 121601	S.-W. Lin <i>et al.</i>	(BELLE Collab.)	AUBERT,BE	04B	PR D70 112006	B. Aubert <i>et al.</i>	(BABAR Collab.)
SATOYAMA	07	PL B647 67	N. Satoyama <i>et al.</i>	(BELLE Collab.)	AUBERT,BE	04A	PR D70 091106	B. Aubert <i>et al.</i>	(BABAR Collab.)
SCHUEMANN	07	PR D75 092002	J. Schuermann <i>et al.</i>	(BELLE Collab.)	CHAO	04	PR D69 111102	Y. Chao <i>et al.</i>	(BELLE Collab.)
TSAI	07	PR D75 111101	Y.-T. Tsai <i>et al.</i>	(BELLE Collab.)	CHAO	04B	PRL 93 191802	Y. Chao <i>et al.</i>	(BELLE Collab.)
URQUIJO	07	PR D75 032001	P. Urquijo <i>et al.</i>	(BELLE Collab.)	CHISTOV	04	PRL 93 051803	R. Chistov <i>et al.</i>	(BELLE Collab.)
WANG	07B	PR D75 092005	C.H. Wang <i>et al.</i>	(BELLE Collab.)	DRUTSKOY	04	PRL 92 051801	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
WANG	07C	PR D76 052004	M.-Z. Wang <i>et al.</i>	(BELLE Collab.)	GARMASH	04	PR D69 022001	A. Garmash <i>et al.</i>	(BELLE Collab.)
XIE	07	PR D75 017101	Q.L. Xie <i>et al.</i>	(BELLE Collab.)	LEE	04	PRL 93 211801	Y.-J. Lee <i>et al.</i>	(BELLE Collab.)
ABE	06	PR D73 051106	K. Abe <i>et al.</i>	(BELLE Collab.)	MAJUMDER	04	PR D70 111103	G. Majumder <i>et al.</i>	(BELLE Collab.)
ABULENCIA	06J	PRL 96 191801	A. Abulencia <i>et al.</i>	(CDF Collab.)	NAKAO	04	PR D69 112001	M. Nakao <i>et al.</i>	(BELLE Collab.)
ACOSTA	06	PRL 96 202001	D. Acosta <i>et al.</i>	(CDF Collab.)	POLUEKTOV	04	PR D70 072003	A. Poluektov <i>et al.</i>	(BELLE Collab.)
AUBERT	06	PR D73 011101	B. Aubert <i>et al.</i>	(BABAR Collab.)	SCHWANDA	04	PRL 93 131803	C. Schwanda <i>et al.</i>	(BELLE Collab.)
AUBERT	06E	PRL 96 052002	B. Aubert <i>et al.</i>	(BABAR Collab.)	WANG	04	PRL 92 131801	M.Z. Wang <i>et al.</i>	(BELLE Collab.)
AUBERT	06F	PR D73 011103	B. Aubert <i>et al.</i>	(BABAR Collab.)	WANG	04A	PR D70 012001	C.H. Wang <i>et al.</i>	(BELLE Collab.)
AUBERT	06J	PR D73 051105	B. Aubert <i>et al.</i>	(BABAR Collab.)	ZANG	04	PR D69 017101	S.L. Zang <i>et al.</i>	(BELLE Collab.)
AUBERT	06K	PR D73 057101	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	03B	PR D67 032003	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT	06N	PR D74 031103	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	03D	PRL 90 131803	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT	06O	PR D74 032003	B. Aubert <i>et al.</i>	(BABAR Collab.)	ADAM	03	PR D67 032001	N.E. Adam <i>et al.</i>	(CLEO Collab.)
AUBERT	06Z	PR D73 111104	B. Aubert <i>et al.</i>	(BABAR Collab.)	ADAM	03B	PR D68 012004	N.E. Adam <i>et al.</i>	(CLEO Collab.)
AUBERT,B	06A	PR D73 112004	B. Aubert <i>et al.</i>	(BABAR Collab.)	ATHAR	03	PR D68 072003	S.B. Athar <i>et al.</i>	(CLEO Collab.)
AUBERT,B	06C	PR D74 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03K	PRL 90 231801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06E	PR D74 011106	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03L	PRL 91 021801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06G	PRL 97 201801	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03M	PRL 91 051801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06H	PRL 97 201802	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03O	PRL 91 071801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06J	PR D73 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03U	PRL 91 221802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06M	PR D74 031102	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03V	PRL 91 171802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06P	PR D74 031105	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03W	PRL 91 161801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06T	PR D74 051102	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	03X	PR D68 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	06U	PR D74 051104	B. Aubert <i>et al.</i>	(BABAR Collab.)	BORNHEIM	03	PR D68 052002	A. Bornheim <i>et al.</i>	(CLEO Collab.)
AUBERT,B	06Y	PR D74 091105	B. Aubert <i>et al.</i>	(BABAR Collab.)	CHEN	03B	PRL 91 201801	K.-F. Chen <i>et al.</i>	(BELLE Collab.)
AUBERT,BE	06A	PR D74 099903 (errat.)	B. Aubert <i>et al.</i>	(BABAR Collab.)	CHOI	03	PRL 91 262001	S.-K. Choi <i>et al.</i>	(BELLE Collab.)
AUBERT,BE	06C	PRL 97 171805	B. Aubert <i>et al.</i>	(BABAR Collab.)	CSORNA	03	PR D67 112002	S.E. Csorna <i>et al.</i>	(CLEO Collab.)
AUBERT,BE	06G	PRL 97 261801	B. Aubert <i>et al.</i>	(BABAR Collab.)	EDWARDS	03	PR D68 011102	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
AUBERT,BE	06H	PRL 97 261803	B. Aubert <i>et al.</i>	(BABAR Collab.)	FANG	03	PRL 90 071801	F. Fang <i>et al.</i>	(BELLE Collab.)
AUBERT,BE	06J	PR D74 111102	B. Aubert <i>et al.</i>	(BABAR Collab.)	HUANG	03	PRL 91 241802	H.-C. Huang <i>et al.</i>	(BELLE Collab.)
AUBERT,BE	06M	PR D74 071101	B. Aubert <i>et al.</i>	(BABAR Collab.)	ISHIKAWA	03	PRL 91 261601	A. Ishikawa <i>et al.</i>	(BELLE Collab.)
CHISTOV	06A	PR D74 111105	R. Chistov <i>et al.</i>	(BELLE Collab.)	KROKOVNY	03B	PRL 91 262002	P. Krokovny <i>et al.</i>	(BELLE Collab.)
FANG	06	PR D74 012007	F. Fang <i>et al.</i>	(BELLE Collab.)	SWAIN	03	PR D68 051101	S.K. Swain <i>et al.</i>	(BELLE Collab.)
GABYSHEV	06	PRL 97 202003	N. Gabyshev <i>et al.</i>	(BELLE Collab.)	UNNO	03	PR D68 011103	Y. Unno <i>et al.</i>	(BELLE Collab.)
GABYSHEV	06A	PRL 97 242001	N. Gabyshev <i>et al.</i>	(BELLE Collab.)	ZHANG	03	PRL 91 221801	J. Zhang <i>et al.</i>	(BELLE Collab.)
GARMASH	06	PRL 96 251801	A. Garmash <i>et al.</i>	(BELLE Collab.)	ABE	02B	PRL 88 021802	K. Abe <i>et al.</i>	(BELLE Collab.)
GOKHROO	06	PRL 97 162002	G. Gokhroo <i>et al.</i>	(BELLE Collab.)	ABE	02B	PRL 88 031802	K. Abe <i>et al.</i>	(BELLE Collab.)
IKADO	06	PRL 97 251802	K. Ikado <i>et al.</i>	(BELLE Collab.)	ABE	02H	PRL 88 171801	K. Abe <i>et al.</i>	(BELLE Collab.)
JEN	06	PR D74 111101	C.-M. Jen <i>et al.</i>	(BELLE Collab.)	ABE	02K	PRL 88 181803	K. Abe <i>et al.</i>	(BELLE Collab.)
KUMAR	06	PR D74 051103	R. Kumar <i>et al.</i>	(BELLE Collab.)	ABE	02N	PL B538 11	K. Abe <i>et al.</i>	(BELLE Collab.)
MOHAPATRA	06	PRL 96 221601	D. Mohapatra <i>et al.</i>	(BELLE Collab.)	ABE	02O	PR D65 091103	K. Abe <i>et al.</i>	(BELLE Collab.)
POLUEKTOV	06	PR D73 112009	A. Poluektov <i>et al.</i>	(BELLE Collab.)	ABE	02W	PRL 89 151802	K. Abe <i>et al.</i>	(BELLE Collab.)
SCHUEMANN	06	PRL 97 061802	J. Schuermann <i>et al.</i>	(BELLE Collab.)	ACOSTA	02C	PR D65 092009	D. Acosta <i>et al.</i>	(CDF Collab.)
SONI	06	PL B634 155	N. Soni <i>et al.</i>	(BELLE Collab.)	ACOSTA	02F	PR D66 052005	D. Acosta <i>et al.</i>	(CDF Collab.)
ABE	05A	PRL 94 221805	K. Abe <i>et al.</i>	(BELLE Collab.)	AHMED	02B	PR D66 031101	S. Ahmed <i>et al.</i>	(CLEO Collab.)
ABE	05B	PR D71 072003	K. Abe <i>et al.</i>	(BELLE Collab.)	AUBERT	02	PR D65 032001	B. Aubert <i>et al.</i>	(BABAR Collab.)
Also		PR D71 079901 (errat.)	K. Abe <i>et al.</i>	(BELLE Collab.)	AUBERT	02C	PRL 88 101805	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE	05G	PRL 95 231802	K. Abe <i>et al.</i>	(BELLE Collab.)	AUBERT	02E	PR D65 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)
ACOSTA	05J	PRL 95 031801	D. Acosta <i>et al.</i>	(CDF Collab.)	AUBERT	02F	PR D65 091101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	05	PRL 94 011801	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	02L	PRL 88 241801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	05B	PR D71 031501	B. Aubert <i>et al.</i>	(BABAR Collab.)	BRIERE	02	PRL 89 081803	R. Briere <i>et al.</i>	(CLEO Collab.)
AUBERT	05G	PR D72 032004	B. Aubert <i>et al.</i>	(BABAR Collab.)	CASEY	02	PR D66 092002	B.C.K. Casey <i>et al.</i>	(BELLE Collab.)
AUBERT	05H	PRL 94 101801	B. Aubert <i>et al.</i>	(BABAR Collab.)	CHEN	02B	PL B546 196	K.-F. Chen <i>et al.</i>	(BELLE Collab.)
AUBERT	05J	PRL 94 141801	B. Aubert <i>et al.</i>	(BABAR Collab.)	DRUTSKOY	02	PL B542 171	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
AUBERT	05K	PRL 94 171801	B. Aubert <i>et al.</i>	(BABAR Collab.)	DYTMAN	02	PR D66 091101	S.A. Dytman <i>et al.</i>	(CLEO Collab.)
AUBERT	05L	PRL 94 181802	B. Aubert <i>et al.</i>	(BABAR Collab.)	ECKHART	02	PRL 89 251801	E. Eckhart <i>et al.</i>	(CLEO Collab.)
AUBERT	05M	PRL 94 191802	B. Aubert <i>et al.</i>	(BABAR Collab.)	EDWARDS	02B	PR D65 111102	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
AUBERT	05N	PR D71 031102	B. Aubert <i>et al.</i>	(BABAR Collab.)	GABYSHEV	02	PR D66 091102	N. Gabyshev <i>et al.</i>	(BELLE Collab.)
AUBERT	05O	PR D71 031103	B. Aubert <i>et al.</i>	(BABAR Collab.)	GARMASH	02	PR D65 092005	A. Garmash <i>et al.</i>	(BELLE Collab.)
AUBERT	05R	PR D71 071103	B. Aubert <i>et al.</i>	(BABAR Collab.)	GODANG	02	PRL 88 021802	R. Godang <i>et al.</i>	(CLEO Collab.)
AUBERT	05U	PR D71 091103	B. Aubert <i>et al.</i>	(BABAR Collab.)	GORDON	02	PL B542 183	A. Gordon <i>et al.</i>	(BELLE Collab.)
AUBERT	05X	PR D71 111101	B. Aubert <i>et al.</i>	(BABAR Collab.)	LU	02	PRL 89 191801	R.-S. Lu <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05B	PRL 95 041804	B. Aubert <i>et al.</i>	(BABAR Collab.)	MAHAPATRA	02	PRL 88 101803	R. Mahapatra <i>et al.</i>	(CLEO Collab.)
AUBERT,B	05E	PR D72 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)	NISHIDA	02	PRL 89 231801	S. Nishida <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05G	PR D72 052002	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	01H	PRL 87 101801	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05K	PRL 95 131803	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	01I	PRL 87 111801	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05L	PR D72 051101	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	01K	PR D64 071101	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05N	PR D72 072003	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	01L	PRL 87 161601	K. Abe <i>et al.</i>	(BELLE Collab.)
Also		PR D72 099903 (errat.)	B. Aubert <i>et al.</i>	(BABAR Collab.)	ABE	01B	PL B517 309	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05O	PR D72 051102	B. Aubert <i>et al.</i>	(BABAR Collab.)	ALEXANDER	01B	PR D64 092001	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
AUBERT,B	05T	PR D72 071102	B. Aubert <i>et al.</i>	(BABAR Collab.)	AMMAR	01B	PRL 87 271801	R. Ammar <i>et al.</i>	(CLEO Collab.)
AUBERT,B	05U	PR D72 071103	B. Aubert <i>et al.</i>	(BABAR Collab.)	ANDERSON	01B	PRL 87 181803	S. Anderson <i>et al.</i>	(CLEO Collab.)
AUBERT,B	05V	PR D72 071104	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	01D	PRL 87 151801	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	05Y	PRL 95 121802	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	01E	PRL 87 151802	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,BE	05E	PRL 95 221801	B. Aubert <i>et al.</i>	(BABAR Collab.)	AUBERT	01F	PRL 87 201803	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHANG	05	PR D71 072007	M.-C. Chang <i>et al.</i>	(BELLE Collab.)	AUBERT	01G	PRL 87 221802	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHANG	05A	PR D71 091106	P. Chang <i>et al.</i>	(BELLE Collab.)	BARATE	01E	EPJ C19 213	R. Barate <i>et al.</i>	(ALEPH Collab.)
CHAO	05A	PR D71 031502	Y. Chao <i>et al.</i>	(BELLE Collab.)	BRIERE	01	PRL 86 3718	R.A. Biere <i>et al.</i>	(CLEO Collab.)
CHEN	05A	PRL 94 221804	K.-F. Chen <i>et al.</i>	(BELLE Collab.)	BROWDER	01	PRL 86 2950	T.E. Browder <i>et al.</i>	(CLEO Collab.)
GARMASH	05	PR D71 092003	A. Garmash <i>et al.</i>	(BELLE Collab.)	EDWARDS	01	PRL 86 30	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ITO	05	PRL 95 091601	R. Itoh <i>et al.</i>	(BELLE Collab.)	A. GRITSAN	01	PR D64 077501	A. Gritsan <i>et al.</i>	(CLEO Collab.)
LEE	05	PRL 95 061802	Y.-J. Lee <i>et al.</i>	(BELLE Collab.)	RICHICHI	01	PR D63 031103	S.J. Richichi <i>et al.</i>	(CLEO Collab.)
LIVENTSEV	05	PR D72 051109	D. Liventsev <i>et al.</i>	(BELLE Collab.)	ABBIENDI	00B	PL B476 233	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
MAJUMDER	05	PRL 95 041803	G. Majumder <i>et al.</i>	(BELLE Collab.)	ABE	00C	PR D62 071101	K. Abe <i>et al.</i>	(SLD Collab.)
MOHAPATRA	05	PR D72 011101	D. Mohapatra <i>et al.</i>	(BELLE Collab.)	AHMED	00B	PR D62 112003	S. Ahmed <i>et al.</i>	(CLEO Collab.)
NISHIDA	05	PL B610 23	S. Nishida <i>et al.</i>	(BELLE Collab.)	ANASTASSOV	00	PRL 84 1393	A. Anastassov <i>et al.</i>	(CLEO Collab.)
OKABE	05	PL B614 27	T. Okabe <i>et al.</i>	(BELLE Collab.)	BARATE	00R	PL B492 275	R. Barate <i>et al.</i>	(ALEPH Collab.)
SAIGO	05	PRL 94							

# Meson Particle Listings

## $B^\pm, B^0$

BARATE	98Q	EPJ C4 387	R. Barate <i>et al.</i>	(ALEPH Collab.)
BEHRENS	98	PRL 80 3710	B.H. Behrens <i>et al.</i>	(CLEO Collab.)
BERGFELD	98	PRL 81 272	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
BRANDENB...	98	PRL 80 2762	G. Brandenbrug <i>et al.</i>	(CLEO Collab.)
CAPRINI	98	NP B530 153	I. Caprini, L. Lellouch, M. Neubert	(BCEP, CERN)
GODANG	98	PRL 80 3456	R. Godang <i>et al.</i>	(CLEO Collab.)
ABE	97J	PRL 79 590	K. Abe <i>et al.</i>	(SLD Collab.)
ACCIARRI	97F	PL B396 327	M. Acciari <i>et al.</i>	(L3 Collab.)
ARTUSO	97	PL B399 321	M. Artuso <i>et al.</i>	(CLEO Collab.)
ATHANAS	97	PRL 79 2208	M. Athanas <i>et al.</i>	(CLEO Collab.)
BROWDER	97	PR D56 11	T. Browder <i>et al.</i>	(CLEO Collab.)
FU	97	PRL 79 3125	X. Fu <i>et al.</i>	(CLEO Collab.)
JESSOP	97	PRL 79 4533	C.P. Jessop <i>et al.</i>	(CLEO Collab.)
ABE	96B	PR D53 3496	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	96C	PRL 76 4462	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	96H	PRL 76 2015	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	96L	PRL 76 4675	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	96Q	PR D54 6596	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	96R	PRL 77 5176	F. Abe <i>et al.</i>	(CDF Collab.)
ADAM	96D	ZPHY C72 207	W. Adam <i>et al.</i>	(DELPHI Collab.)
ALEXANDER	96T	PRL 77 5000	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
ASNER	96	PR D53 1039	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BARISH	96B	PRL 76 1570	B.C. Barish <i>et al.</i>	(CLEO Collab.)
BERGFELD	96B	PRL 77 4503	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
BISHAI	96	PL B369 186	M. Bishai <i>et al.</i>	(CLEO Collab.)
BUSKULIC	96J	ZPHY C71 31	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
GIBAUT	96	PR D53 4734	D. Gibaut <i>et al.</i>	(CLEO Collab.)
PDG	96	PR D54 1	R. M. Barnett <i>et al.</i>	(PDG Collab.)
ABREU	95N	PL B357 255	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95Q	ZPHY C68 13	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ADAM	95T	ZPHY C68 363	W. Adam <i>et al.</i>	(DELPHI Collab.)
AKERS	95T	ZPHY C67 379	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	95D	PL B353 554	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	95	PL B341 435	J. Alexander <i>et al.</i>	(CLEO Collab.)
Also		PL B347 469 (erratum)	M. Artuso <i>et al.</i>	(CLEO Collab.)
ARTUSO	95	PRL 75 785	B.C. Barish <i>et al.</i>	(CLEO Collab.)
BARISH	95	PR D51 1014	B.C. Barish <i>et al.</i>	(CLEO Collab.)
BUSKULIC	95	PL B343 444	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
ABE	94D	PRL 72 3456	F. Abe <i>et al.</i>	(CDF Collab.)
ALAM	94	PR D50 43	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	94D	PL B355 526	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ATHANAS	94	PRL 73 3503	M. Athanas <i>et al.</i>	(CLEO Collab.)
Also		PRL 74 3090 (erratum)	M. Athanas <i>et al.</i>	(CLEO Collab.)
PDG	94	PR D50 1173	L. Montanet <i>et al.</i>	(CERN, LBL, BOST+)
STONE	94	HEPSY 93-11	S. Stone	
Published in B		Decays, 2nd Edition, World Scientific, Singapore		
ABREU	93D	ZPHY C57 181	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	93G	PL B312 253	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTION	93C	PL B307 247	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ALBRECHT	93E	ZPHY C60 11	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	93B	PL B319 365	J. Alexander <i>et al.</i>	(CLEO Collab.)
AMMAR	93	PRL 71 674	R. Ammar <i>et al.</i>	(CLEO Collab.)
BEAN	93B	PRL 70 2681	A. Bean <i>et al.</i>	(CLEO Collab.)
BUSKULIC	93D	PL B307 194	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
Also		PL B325 537 (erratum)	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
SANGHERA	93	PR D47 791	S. Sanghera <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92C	PL B275 195	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92E	PL B277 209	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92G	ZPHY C54 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BORTOLETTO	92	PR D45 21	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
BUSKULIC	92G	PL B295 396	D. Buskalic <i>et al.</i>	(BCEP Collab.)
ALBRECHT	91B	PL B254 288	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	91C	PL B255 297	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	91E	PL B262 148	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BERKELMAN	91	ARNPS 41 1	K. Berkelman, S. Stone	(CORN, SYRA)
"Decays of B Mesons"				
FULTON	91	PR D43 651	R. Fulton <i>et al.</i>	(CLEO Collab.)
ALBRECHT	90B	PL B241 278	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	90J	ZPHY C48 543	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANTREASYAN	90B	ZPHY C48 553	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
BORTOLETTO	90	PRL 64 2117	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
Also		PR D45 21	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
WEIR	90B	PR D41 1384	A.J. Weir <i>et al.</i>	(Mark II Collab.)
ALBRECHT	89G	PL B229 304	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AVERY	89B	PL B223 470	P. Avery <i>et al.</i>	(CLEO Collab.)
BEBEK	89	PRL 62 8	C. Bebek <i>et al.</i>	(CLEO Collab.)
BORTOLETTO	89	PRL 62 2436	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
ALBRECHT	88F	PL B209 119	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88K	PL B215 424	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87C	PL B185 218	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87D	PL B199 451	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AVERY	87	PL B183 429	P. Avery <i>et al.</i>	(CLEO Collab.)
BEBEK	87	PR D36 1289	C. Bebek <i>et al.</i>	(CLEO Collab.)
ALAM	86	PR D34 3279	M.S. Alam <i>et al.</i>	(CLEO Collab.)
PDG	86	PL 170B 1	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)
GILES	84	PR D30 2279	R. Giles <i>et al.</i>	(CLEO Collab.)

### $B^0$

Quantum numbers not measured. Values shown are quark-model predictions.

See also the  $B^\pm/B^0$  ADMIXTURE and  $B^\pm/B^0/B_S^0/b$ -baryon ADMIXTURE sections.

See the Note "Production and Decay of  $b$ -flavored Hadrons" at the beginning of the  $B^\pm$  Particle Listings and the Note on " $B^0$ - $\bar{B}^0$  Mixing" near the end of the  $B^0$  Particle Listings.

### $B^0$ MASS

The fit uses  $m_{B^{+}}$ , ( $m_{B^0} - m_{B^{+}}$ ), and  $m_{B^0}$  to determine  $m_{B^{+}}$ ,  $m_{B^0}$ , and the mass difference.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5279.62±0.15 OUR FIT</b>		Error includes scale factor of 1.1.		
<b>5279.55±0.26 OUR AVERAGE</b>				
5279.6 ±0.2 ±0.0	1	AAD	13u	ATLS $pp$ at 7 TeV
5279.58±0.15±0.28	2	AAIJ	12E	LHCB $pp$ at 7 TeV
5279.63±0.53±0.33	3	ACOSTA	06	CDF $p\bar{p}$ at 1.96 TeV
5279.1 ±0.7 ±0.3	135	4	CSORNA	00 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
5281.3 ±2.2 ±1.4	51	ABE	96B	CDF $p\bar{p}$ at 1.8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

5279.2 ±0.54±2.0	340	ALAM	94	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
5278.0 ±0.4 ±2.0		BORTOLETTO	92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
5279.6 ±0.7 ±2.0	40	5	ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
5278.2 ±1.0 ±3.0	40	4	ALBRECHT	87c	ARG $e^+e^- \rightarrow \Upsilon(4S)$
5279.5 ±1.6 ±3.0	7	6	ALBRECHT	87D	ARG $e^+e^- \rightarrow \Upsilon(4S)$
5280.6 ±0.8 ±2.0		BEBEK	87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

- 1 Measured with  $B_D^0 \rightarrow J/\psi(\mu^+\mu^-) K_S^0(\pi^+\pi^-)$  decays.
- 2 Uses  $B^0 \rightarrow J/\psi K^0$  fully reconstructed decays.
- 3 Uses exclusively reconstructed final states containing a  $J/\psi \rightarrow \mu^+\mu^-$  decays.
- 4 CSORNA 00 uses fully reconstructed 135  $B^0 \rightarrow J/\psi(\ell) K_S^0$  events and invariant masses without beam constraint.
- 5 ALBRECHT 90J assumes 10580 for  $\Upsilon(4S)$  mass. Supersedes ALBRECHT 87c and ALBRECHT 87D.
- 6 Found using fully reconstructed decays with  $J/\psi$ . ALBRECHT 87D assume  $m_{\Upsilon(4S)} = 10577$  MeV.

### $m_{B^0} - m_{B^+}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>0.31±0.06 OUR FIT</b>			
<b>0.32±0.05 OUR AVERAGE</b>			
0.20±0.17±0.11	1	AAIJ	12E LHCB $pp$ at 7 TeV
0.33±0.05±0.03	2	AUBERT	08AF BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.53±0.67±0.14	3	ACOSTA	06 CDF $p\bar{p}$ at 1.96 TeV
0.41±0.25±0.19		ALAM	94 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
−0.4 ±0.6 ±0.5		BORTOLETTO	92 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
−0.9 ±1.2 ±0.5		ALBRECHT	90J ARG $e^+e^- \rightarrow \Upsilon(4S)$
2.0 ±1.1 ±0.3	4	BEBEK	87 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
1 Uses exclusively reconstructed final states containing a $J/\psi \rightarrow \mu^+\mu^-$ decay.			
2 Uses the $B$ -momentum distributions in the $e^+e^-$ rest frame.			
3 Uses exclusively reconstructed final states containing a $J/\psi \rightarrow \mu^+\mu^-$ decays.			
4 BEBEK 87 actually measure the difference between half of $E_{cm}$ and the $B^\pm$ or $B^0$ mass, so the $m_{B^0} - m_{B^\pm}$ is more accurate. Assume $m_{\Upsilon(4S)} = 10580$ MeV.			

### $m_{B_H^0} - m_{B_L^0}$

See the  $B^0$ - $\bar{B}^0$  MIXING PARAMETERS section near the end of these  $B^0$  Listings.

### $B^0$ MEAN LIFE

See  $B^\pm/B^0/B_S^0/b$ -baryon ADMIXTURE section for data on  $B$ -hadron mean life averaged over species of bottom particles.

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

VALUE (10 <sup>−12</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
1.520±0.004 OUR EVALUATION				
1.534±0.019±0.021		<sup>1</sup> ABAZOV	15A D0	$p\bar{p}$ at 1.96 TeV
1.499±0.013±0.005		<sup>2</sup> AAIJ	14E LHCB	$pp$ at 7 TeV
1.524±0.006±0.004		<sup>3</sup> AAIJ	14E LHCB	$pp$ at 7 TeV
1.524±0.011±0.004		<sup>4</sup> AAIJ	14R LHCB	$pp$ at 7 TeV
1.509±0.012±0.018		<sup>5</sup> AAD	13U ATLS	$pp$ at 7 TeV
1.508±0.025±0.043		<sup>2</sup> ABAZOV	12U D0	$p\bar{p}$ at 1.96 TeV
1.507±0.010±0.008		<sup>6</sup> AALTONEN	11 CDF	$p\bar{p}$ at 1.96 TeV
1.414±0.018±0.034		<sup>7</sup> ABAZOV	09E D0	$p\bar{p}$ at 1.96 TeV
1.504±0.013 <sup>+0.018</sup> <sub>−0.013</sub>		<sup>8</sup> AUBERT	06G BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.534±0.008±0.010		<sup>9</sup> ABE	05B BELL	$e^+e^- \rightarrow \Upsilon(4S)$
1.531±0.021±0.031		<sup>10</sup> ABDALLAH	04E DLPH	$e^+e^- \rightarrow Z$
1.523 <sup>+0.024</sup> <sub>−0.023</sub> ±0.022		<sup>11</sup> AUBERT	03C BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.533±0.034±0.038		<sup>12</sup> AUBERT	03H BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.497±0.073±0.032		<sup>13</sup> ACOSTA	02C CDF	$p\bar{p}$ at 1.8 TeV
1.529±0.012±0.029		<sup>14</sup> AUBERT	02H BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.546±0.032±0.022		<sup>15</sup> AUBERT	01F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.541±0.028±0.023		<sup>14</sup> ABBIENDI,G	00B OPAL	$e^+e^- \rightarrow Z$
1.518±0.053±0.034		<sup>16</sup> BARATE	00R ALEP	$e^+e^- \rightarrow Z$
1.523±0.057±0.053		<sup>17</sup> ABBIENDI	99J OPAL	$e^+e^- \rightarrow Z$
1.474±0.039 <sup>+0.052</sup> <sub>−0.051</sub>		<sup>16</sup> ABE	98Q CDF	$p\bar{p}$ at 1.8 TeV
1.52 ±0.06 ±0.04		<sup>17</sup> ACCIARRI	98S L3	$e^+e^- \rightarrow Z$
1.64 ±0.08 ±0.08		<sup>17</sup> ABE	97J SLD	$e^+e^- \rightarrow Z$
1.532±0.041±0.040		<sup>18</sup> ABREU	97F DLPH	$e^+e^- \rightarrow Z$
1.25 <sup>+0.15</sup> <sub>−0.13</sub> ±0.05	121	<sup>13</sup> BUSKULIC	96J ALEP	$e^+e^- \rightarrow Z$
1.49 <sup>+0.17</sup> <sub>−0.15</sub> <sup>+0.08</sup> <sub>−0.06</sub>		<sup>19</sup> BUSKULIC	96J ALEP	$e^+e^- \rightarrow Z$
1.61 <sup>+0.14</sup> <sub>−0.13</sub> ±0.08		<sup>16,20</sup> ABREU	95Q DLPH	$e^+e^- \rightarrow Z$
1.63 ±0.14 ±0.13		<sup>21</sup> ADAM	95Q DLPH	$e^+e^- \rightarrow Z$
1.53 ±0.12 ±0.08		<sup>16,22</sup> AKERS	95T OPAL	$e^+e^- \rightarrow Z$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.501^{+0.078}_{-0.074} \pm 0.050$	<sup>2</sup> ABAZOV	07s	D0	Repl. by ABAZOV 12u
$1.524 \pm 0.030 \pm 0.016$	<sup>2</sup> ABULENCIA	07A	CDF	Repl. by AALTONEN 11
$1.473^{+0.052}_{-0.050} \pm 0.023$	<sup>7</sup> ABAZOV	05B	D0	Repl. by ABAZOV 05w
$1.40^{+0.11}_{-0.10} \pm 0.03$	<sup>2</sup> ABAZOV	05c	D0	Repl. by ABAZOV 07s
$1.530 \pm 0.043 \pm 0.023$	<sup>7</sup> ABAZOV	05W	D0	Repl. by ABAZOV 09E
$1.54 \pm 0.05 \pm 0.02$	<sup>23</sup> ACOSTA	05	CDF	Repl. by AALTONEN 11
$1.554 \pm 0.030 \pm 0.019$	<sup>15</sup> ABE	02H	BELL	Repl. by ABE 05B
$1.58 \pm 0.09 \pm 0.02$	<sup>13</sup> ABE	98B	CDF	Repl. by ACOSTA 02c
$1.54 \pm 0.08 \pm 0.06$	<sup>16</sup> ABE	96c	CDF	Repl. by ABE 98Q
$1.55 \pm 0.06 \pm 0.03$	<sup>24</sup> BUSKULIC	96J	ALEP	$e^+e^- \rightarrow Z$
$1.61 \pm 0.07 \pm 0.04$	<sup>16</sup> BUSKULIC	96J	ALEP	Repl. by BARATE 00R
$1.62 \pm 0.12$	<sup>25</sup> ADAM	95	DLPH	$e^+e^- \rightarrow Z$
$1.57 \pm 0.18 \pm 0.08$	<sup>121</sup> <sup>13</sup> ABE	94D	CDF	Repl. by ABE 98B
$1.17^{+0.29}_{-0.23} \pm 0.16$	<sup>96</sup> <sup>16</sup> ABREU	93D	DLPH	Sup. by ABREU 95Q
$1.55 \pm 0.25 \pm 0.18$	<sup>76</sup> <sup>21</sup> ABREU	93G	DLPH	Sup. by ADAM 95
$1.51^{+0.24}_{-0.23} \pm 0.14$	<sup>78</sup> <sup>16</sup> ACTON	93c	OPAL	Sup. by AKERS 95T
$1.52^{+0.20}_{-0.18} \pm 0.07$	<sup>77</sup> <sup>16</sup> BUSKULIC	93D	ALEP	Sup. by BUSKULIC 96J
$1.20^{+0.52}_{-0.36} \pm 0.14$	<sup>15</sup> <sup>26</sup> WAGNER	90	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
$0.82^{+0.57}_{-0.37} \pm 0.27$	<sup>27</sup> AVERILL	89	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

- Measured using  $B^0 \rightarrow D^{*-}\mu^+\nu X$  decays.
- Measured mean life using  $B^0 \rightarrow J/\psi K_S^0$  decays.
- Measured using  $B^0 \rightarrow J/\psi K^{*0}$  decays.
- Measured using  $B^0 \rightarrow K^+\pi^-$  decays.
- Measured with  $B_d^0 \rightarrow J/\psi(\mu^+\mu^-) K_S^0(\pi^+\pi^-)$  decays.
- Measured mean life using fully reconstructed decays ( $J/\psi K^{*0}$ ).
- Measured mean life using  $B^0 \rightarrow J/\psi K^{*0}$  decays.
- Measured using a simultaneous fit of the  $B^0$  lifetime and  $\bar{B}^0 B^0$  oscillation frequency  $\Delta m_d$  in the partially reconstructed  $B^0 \rightarrow D^{*-}\ell\nu$  decays.
- Measurement performed using a combined fit of  $CP$ -violation, mixing and lifetimes.
- Measurement performed using an inclusive reconstruction and  $B$  flavor identification technique.
- AUBERT 03c uses a sample of approximately 14,000 exclusively reconstructed  $B^0 \rightarrow D^*(2010)^-\ell\nu$  and simultaneously measures the lifetime and oscillation frequency.
- Measurement performed with decays  $B^0 \rightarrow D^{*-}\pi^+$  and  $B^0 \rightarrow D^{*-}\rho^+$  using a partial reconstruction technique.
- Measured mean life using fully reconstructed decays.
- Data analyzed using partially reconstructed  $\bar{B}^0 \rightarrow D^{*+}\ell^-\bar{\nu}$  decays.
- Events are selected in which one  $B$  meson is fully reconstructed while the second  $B$  meson is reconstructed inclusively.
- Data analyzed using  $D/D^*\ell X$  event vertices.
- Data analyzed using charge of secondary vertex.
- Data analyzed using inclusive  $D/D^*\ell X$ .
- Measured mean life using partially reconstructed  $D^{*-}\pi^+ X$  vertices.
- ABREU 95Q assumes  $B(B^0 \rightarrow D^{*-}\ell^+\nu_\ell) = 3.2 \pm 1.7\%$ .
- Data analyzed using vertex-charge technique to tag  $B$  charge.
- AKERS 95T assumes  $B(B^0 \rightarrow D_s^{*0} D^0(*) = 5.0 \pm 0.9\%$  to find  $B^+/B^0$  yield.
- Measured using the time-dependent angular analysis of  $B_d^0 \rightarrow J/\psi K^{*0}$  decays.
- Combined result of  $D/D^*\ell X$  analysis, fully reconstructed  $B$  analysis, and partially reconstructed  $D^{*-}\pi^+ X$  analysis.
- Combined ABREU 95Q and ADAM 95 result.
- WAGNER 90 tagged  $B^0$  mesons by their decays into  $D^{*-}e^+\nu$  and  $D^{*-}\mu^+\nu$  where the  $D^{*-}$  is tagged by its decay into  $\pi^-\bar{D}^0$ .
- AVERILL 89 is an estimate of the  $B^0$  mean lifetime assuming that  $B^0 \rightarrow D^{*+} + X$  always.

 $\tau_{B^0}/\tau_{B^+}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.000 \pm 0.008 \pm 0.009</math></b>	<sup>1</sup> AAIJ	14E	LHCB $pp$ at 7 TeV

- Measured using  $B^0 \rightarrow J/\psi K^{*0}$  decays.

MEAN LIFE RATIO  $\tau_{B^+}/\tau_{B^0}$  $\tau_{B^+}/\tau_{B^0}$  (direct measurements)

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.076 \pm 0.004</math> OUR EVALUATION</b>				
$1.074 \pm 0.005 \pm 0.003$	<sup>1</sup> AAIJ	14E	LHCB	$pp$ at 7 TeV
$1.088 \pm 0.009 \pm 0.004$	<sup>2</sup> AALTONEN	11	CDF	$p\bar{p}$ at 1.96 TeV
$1.080 \pm 0.016 \pm 0.014$	<sup>3</sup> ABZOV	05D	D0	$p\bar{p}$ at 1.96 TeV
$1.066 \pm 0.008 \pm 0.008$	<sup>4</sup> ABE	05B	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.060 \pm 0.021 \pm 0.024$	<sup>5</sup> ABDALLAH	04E	DLPH	$e^+e^- \rightarrow Z$
$1.093 \pm 0.066 \pm 0.028$	<sup>6</sup> ACOSTA	02c	CDF	$p\bar{p}$ at 1.8 TeV
$1.082 \pm 0.026 \pm 0.012$	<sup>7</sup> AUBERT	01F	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.085 \pm 0.059 \pm 0.018$	<sup>3</sup> BARATE	00R	ALEP	$e^+e^- \rightarrow Z$

$1.079 \pm 0.064 \pm 0.041$	<sup>8</sup> ABBIENDI	99J	OPAL	$e^+e^- \rightarrow Z$
$1.110 \pm 0.056^{+0.033}_{-0.030}$	<sup>3</sup> ABE	98Q	CDF	$p\bar{p}$ at 1.8 TeV
$1.09 \pm 0.07 \pm 0.03$	<sup>8</sup> ACCIARRI	98S	L3	$e^+e^- \rightarrow Z$
$1.01 \pm 0.07 \pm 0.06$	<sup>8</sup> ABE	97J	SLD	$e^+e^- \rightarrow Z$
$1.27^{+0.23}_{-0.19} \pm 0.03$	<sup>6</sup> BUSKULIC	96J	ALEP	$e^+e^- \rightarrow Z$
$1.00^{+0.17}_{-0.15} \pm 0.10$	<sup>3,9</sup> ABREU	95Q	DLPH	$e^+e^- \rightarrow Z$
$1.06^{+0.13}_{-0.11} \pm 0.10$	<sup>10</sup> ADAM	95	DLPH	$e^+e^- \rightarrow Z$
$0.99 \pm 0.14^{+0.05}_{-0.04}$	<sup>3,11</sup> AKERS	95T	OPAL	$e^+e^- \rightarrow Z$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.091 \pm 0.023 \pm 0.014$	<sup>7</sup> ABE	02H	BELL	Repl. by ABE 05B
$1.06 \pm 0.07 \pm 0.02$	<sup>6</sup> ABE	98B	CDF	Repl. by ACOSTA 02c
$1.01 \pm 0.11 \pm 0.02$	<sup>3</sup> ABE	96c	CDF	Repl. by ABE 98Q
$1.03 \pm 0.08 \pm 0.02$	<sup>12</sup> BUSKULIC	96J	ALEP	$e^+e^- \rightarrow Z$
$0.98 \pm 0.08 \pm 0.03$	<sup>3</sup> BUSKULIC	96J	ALEP	Repl. by BARATE 00R
$1.02 \pm 0.16 \pm 0.05$	<sup>269</sup> <sup>6</sup> ABE	94D	CDF	Repl. by ABE 98B
$1.11^{+0.51}_{-0.39} \pm 0.11$	<sup>188</sup> <sup>3</sup> ABREU	93D	DLPH	Sup. by ABREU 95Q
$1.01^{+0.29}_{-0.22} \pm 0.12$	<sup>253</sup> <sup>10</sup> ABREU	93G	DLPH	Sup. by ADAM 95
$1.0^{+0.33}_{-0.25} \pm 0.08$	<sup>130</sup> ACTON	93c	OPAL	Sup. by AKERS 95T
$0.96^{+0.19}_{-0.15} \pm 0.18$	<sup>154</sup> <sup>3</sup> BUSKULIC	93D	ALEP	Sup. by BUSKULIC 96J

- Measured using  $B \rightarrow J/\psi K^{(*)}$  decays.

- Measured mean life using fully reconstructed decays ( $J/\psi K^{(*)}$ ).

- Data analyzed using  $D/D^*\mu X$  vertices.

- Measurement performed using a combined fit of  $CP$ -violation, mixing and lifetimes.

- Measurement performed using an inclusive reconstruction and  $B$  flavor identification technique.

- Measured using fully reconstructed decays.

- Events are selected in which one  $B$  meson is fully reconstructed while the second  $B$  meson is reconstructed inclusively.

- Data analyzed using charge of secondary vertex.

- ABREU 95Q assumes  $B(B^0 \rightarrow D^{*-}\ell^+\nu_\ell) = 3.2 \pm 1.7\%$ .

- Data analyzed using vertex-charge technique to tag  $B$  charge.

- AKERS 95T assumes  $B(B^0 \rightarrow D_s^{*0} D^0(*) = 5.0 \pm 0.9\%$  to find  $B^+/B^0$  yield.

- Combined result of  $D/D^*\ell X$  analysis and fully reconstructed  $B$  analysis.

 $\tau_{B^+}/\tau_{B^0}$  (inferred from branching fractions)

These measurements are inferred from the branching fractions for semileptonic decay or other spectator-dominated decays by assuming that the rates for such decays are equal for  $B^0$  and  $B^+$ . We do not use measurements which assume equal production of  $B^0$  and  $B^+$  because of the large uncertainty in the production ratio.

“OUR EVALUATION” has been obtained by the Heavy Flavor Averaging Group (HFAG) by taking into account correlations between measurements.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.076 \pm 0.034</math> OUR EVALUATION</b>					
<b><math>1.07 \pm 0.04</math> OUR AVERAGE</b>					
$1.07 \pm 0.04 \pm 0.03$			URQUIJO	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$1.067 \pm 0.041 \pm 0.033$			AUBERT,B	06Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.95^{+0.117}_{-0.080} \pm 0.091$	<sup>1</sup> ARTUSO	97	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$1.15 \pm 0.17 \pm 0.06$	<sup>2</sup> JESSOP	97	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$0.93 \pm 0.18 \pm 0.12$	<sup>3</sup> ATHANAS	94	CLE2	Sup. by ARTUSO 97
$0.91 \pm 0.27 \pm 0.21$	<sup>4</sup> ALBRECHT	92c	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$1.0 \pm 0.4$	<sup>4,5</sup> ALBRECHT	92G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$0.89 \pm 0.19 \pm 0.13$	<sup>4</sup> FULTON	91	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
$1.00 \pm 0.23 \pm 0.14$	<sup>4</sup> ALBRECHT	89L	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$0.49 \text{ to } 2.3$	<sup>6</sup> BEAN	87B	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

- ARTUSO 97 uses partial reconstruction of  $B \rightarrow D^{*}\ell\nu_\ell$  and independent of  $B^0$  and  $B^+$  production fraction.

- Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

- ATHANAS 94 uses events tagged by fully reconstructed  $B^-$  decays and partially or fully reconstructed  $B^0$  decays.

- Assumes equal production of  $B^0$  and  $B^+$ .

- ALBRECHT 92G data analyzed using  $B \rightarrow D_s\bar{D}, D_s\bar{D}^*, D_s^*\bar{D}, D_s^*\bar{D}^*$  events.

- BEAN 87B assume the fraction of  $B^0\bar{B}^0$  events at the  $\Upsilon(4S)$  is 0.41.

 $\text{sgn}(\text{Re}(\lambda_{CP})) \Delta\Gamma_{B_d^0} / \Gamma_{B_d^0}$ 

$\Gamma_{B_d^0}$  and  $\Delta\Gamma_{B_d^0}$  are the decay rate average and difference between two  $B_d^0$   $CP$  eigenstates (light – heavy). The  $\lambda_{CP}$  characterizes  $B^0$  and  $\bar{B}^0$  decays to states of charmonium plus  $K_S^0$ , see the review on “ $CP$  Violation” in the reviews section.

“OUR EVALUATION” has been obtained by the Heavy Flavor Averaging Group (HFAG) by taking into account correlations between measurements.

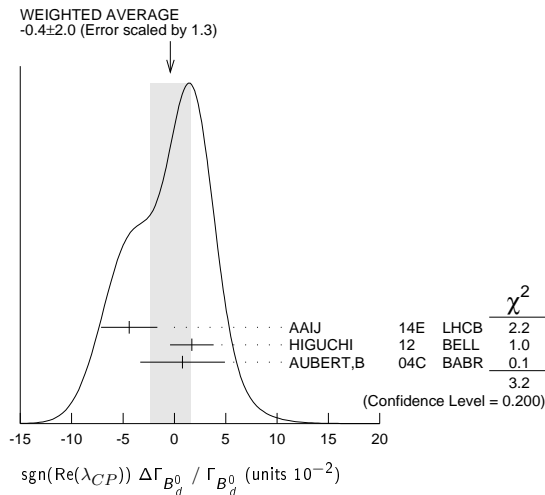
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>-0.3 \pm 1.5</math> OUR EVALUATION</b>				
<b><math>-0.4 \pm 2.0</math> OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
$-4.4 \pm 2.5 \pm 1.1$		<sup>1</sup> AAIJ	14E	LHCB $pp$ at 7 TeV
$1.7 \pm 1.8 \pm 1.1$		<sup>2</sup> HIGUCHI	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.8 \pm 3.7 \pm 1.8$		<sup>3</sup> AUBERT,B	04C	BABR $e^+e^- \rightarrow \Upsilon(4S)$

# Meson Particle Listings

## $B^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.50 \pm 1.38$	ABAZOV	14	D0	$p\bar{p}$ at 1.96 TeV
< 18	95	<sup>4</sup> ABDALLAH	03B	DLPH $e^+e^- \rightarrow Z$
< 80	95	<sup>5</sup> BEHRENS	00B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Measured using the effective lifetimes of $B^0 \rightarrow J/\psi K_S^0$ and $B^0 \rightarrow J/\psi K^{*0}$ decays.				
<sup>2</sup> Reports $-\Delta\Gamma_d/\Gamma_d$ using $B^0 \rightarrow J/\psi K_S^0$ , $J/\psi K_L^0$ , $D^-\pi^+$ , $D^{*-}\pi^+$ , $D^{*-}\rho^+$ , and $D^{*-}\ell^+\nu$ decays.				
<sup>3</sup> Corresponds to 90% confidence range $[-0.084, 0.068]$ .				
<sup>4</sup> Using the measured $\tau_{B^0}=1.55 \pm 0.03$ ps.				
<sup>5</sup> BEHRENS 00B uses high-momentum lepton tags and partially reconstructed $\bar{B}^0 \rightarrow D^{*+}\pi^-$ , $\rho^-$ decays to determine the flavor of the $B$ meson. Assumes $\Delta_{md}=0.478 \pm 0.018$ ps <sup>-1</sup> and $\tau_{B^0}=1.548 \pm 0.032$ ps.				



### $B^0$ DECAY MODES

$\bar{B}^0$  modes are charge conjugates of the modes below. Reactions indicate the weak decay vertex and do not include mixing. Modes which do not identify the charge state of the  $B$  are listed in the  $B^\pm/B^0$  ADMIXTURE section.

The branching fractions listed below assume 50%  $B^0\bar{B}^0$  and 50%  $B^+B^-$  production at the  $\Upsilon(4S)$ . We have attempted to bring older measurements up to date by rescaling their assumed  $\Upsilon(4S)$  production ratio to 50:50 and their assumed  $D$ ,  $D_s$ ,  $D^*$ , and  $\psi$  branching ratios to current values whenever this would affect our averages and best limits significantly.

Indentation is used to indicate a subchannel of a previous reaction. All resonant subchannels have been corrected for resonance branching fractions to the final state so the sum of the subchannel branching fractions can exceed that of the final state.

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm$  anything, the values usually are multiplicities, not branching fractions. They can be greater than one.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $\ell^+\nu_\ell$ anything	[a] ( 10.33 ± 0.28 ) %	
$\Gamma_2$ $e^+\nu_e X_c$	( 10.1 ± 0.4 ) %	
$\Gamma_3$ $D\ell^+\nu_\ell$ anything	( 9.2 ± 0.8 ) %	
$\Gamma_4$ $D^-\ell^+\nu_\ell$	[a] ( 2.19 ± 0.12 ) %	
$\Gamma_5$ $D^-\tau^+\nu_\tau$	( 1.03 ± 0.22 ) %	
$\Gamma_6$ $D^*(2010)^-\ell^+\nu_\ell$	[a] ( 4.93 ± 0.11 ) %	
$\Gamma_7$ $D^*(2010)^-\tau^+\nu_\tau$	( 1.78 ± 0.17 ) %	S=1.1
$\Gamma_8$ $\bar{D}^0\pi^-\ell^+\nu_\ell$	( 4.3 ± 0.6 ) × 10 <sup>-3</sup>	
$\Gamma_9$ $D_0^*(2400)^-\ell^+\nu_\ell$ , $D_0^{*-} \rightarrow \bar{D}^0\pi^-$	( 3.0 ± 1.2 ) × 10 <sup>-3</sup>	S=1.8
$\Gamma_{10}$ $D_2^*(2460)^-\ell^+\nu_\ell$ , $D_2^{*-} \rightarrow \bar{D}^0\pi^-$	( 1.21 ± 0.33 ) × 10 <sup>-3</sup>	S=1.8
$\Gamma_{11}$ $\bar{D}^{(*)}\eta\pi\ell^+\nu_\ell$ ( $\eta \geq 1$ )	( 2.3 ± 0.5 ) %	
$\Gamma_{12}$ $\bar{D}^{*0}\pi^-\ell^+\nu_\ell$	( 4.9 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{13}$ $D_1(2420)^-\ell^+\nu_\ell$ , $D_1^- \rightarrow \bar{D}^{*0}\pi^-$	( 2.80 ± 0.28 ) × 10 <sup>-3</sup>	
$\Gamma_{14}$ $D_1'(2430)^-\ell^+\nu_\ell$ , $D_1'^- \rightarrow \bar{D}^{*0}\pi^-$	( 3.1 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{15}$ $D_2^*(2460)^-\ell^+\nu_\ell$ , $D_2^{*-} \rightarrow \bar{D}^{*0}\pi^-$	( 6.8 ± 1.2 ) × 10 <sup>-4</sup>	

$\Gamma_{16}$ $D^-\pi^+\pi^-\ell^+\nu_\ell$	( 1.3 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{17}$ $D^{*-}\pi^+\pi^-\ell^+\nu_\ell$	( 1.4 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{18}$ $\rho^-\ell^+\nu_\ell$	[a] ( 2.94 ± 0.21 ) × 10 <sup>-4</sup>	
$\Gamma_{19}$ $\pi^-\ell^+\nu_\ell$	[a] ( 1.45 ± 0.05 ) × 10 <sup>-4</sup>	
$\Gamma_{20}$ $\pi^-\mu^+\nu_\mu$		
$\Gamma_{21}$ $\pi^-\tau^+\nu_\tau$	< 2.5	× 10 <sup>-4</sup> CL=90%

### Inclusive modes

$\Gamma_{22}$ $K^\pm$ anything	( 78 ± 8 ) %	
$\Gamma_{23}$ $D^0 X$	( 8.1 ± 1.5 ) %	
$\Gamma_{24}$ $\bar{D}^0 X$	( 47.4 ± 2.8 ) %	
$\Gamma_{25}$ $D^+ X$	< 3.9	% CL=90%
$\Gamma_{26}$ $D^- X$	( 36.9 ± 3.3 ) %	
$\Gamma_{27}$ $D_s^+ X$	( 10.3 ± 2.1 ) %	
$\Gamma_{28}$ $D_s^- X$	< 2.6	% CL=90%
$\Gamma_{29}$ $\Lambda_c^+ X$	< 3.1	% CL=90%
$\Gamma_{30}$ $\bar{\Lambda}_c^- X$	( 5.0 ± 2.1 ) %	
$\Gamma_{31}$ $\bar{c} X$	( 95 ± 5 ) %	
$\Gamma_{32}$ $c X$	( 24.6 ± 3.1 ) %	
$\Gamma_{33}$ $\bar{c} c X$	( 119 ± 6 ) %	

### $D$ , $D^*$ , or $D_s$ modes

$\Gamma_{34}$ $D^-\pi^+$	( 2.52 ± 0.13 ) × 10 <sup>-3</sup>	S=1.1
$\Gamma_{35}$ $D^-\rho^+$	( 7.5 ± 1.2 ) × 10 <sup>-3</sup>	
$\Gamma_{36}$ $D^-\bar{K}^0\pi^+$	( 4.9 ± 0.9 ) × 10 <sup>-4</sup>	
$\Gamma_{37}$ $D^-\bar{K}^*(892)^+$	( 4.5 ± 0.7 ) × 10 <sup>-4</sup>	
$\Gamma_{38}$ $D^-\omega\pi^+$	( 2.8 ± 0.6 ) × 10 <sup>-3</sup>	
$\Gamma_{39}$ $D^-\bar{K}^+$	( 1.86 ± 0.20 ) × 10 <sup>-4</sup>	
$\Gamma_{40}$ $D^-\bar{K}^+\pi^+\pi^-$	( 3.5 ± 0.8 ) × 10 <sup>-4</sup>	
$\Gamma_{41}$ $D^-\bar{K}^+\bar{K}^0$	< 3.1	× 10 <sup>-4</sup> CL=90%
$\Gamma_{42}$ $D^-\bar{K}^+\bar{K}^*(892)^0$	( 8.8 ± 1.9 ) × 10 <sup>-4</sup>	
$\Gamma_{43}$ $\bar{D}^0\pi^+\pi^-$	( 8.8 ± 0.5 ) × 10 <sup>-4</sup>	
$\Gamma_{44}$ $D^*(2010)^-\pi^+$	( 2.74 ± 0.13 ) × 10 <sup>-3</sup>	
$\Gamma_{45}$ $\bar{D}^0\bar{K}^+K^-$	( 4.9 ± 1.2 ) × 10 <sup>-5</sup>	
$\Gamma_{46}$ $D^-\pi^+\pi^+\pi^-$	( 6.0 ± 0.7 ) × 10 <sup>-3</sup>	S=1.1
$\Gamma_{47}$ $(D^-\pi^+\pi^+\pi^-)$ nonresonant	( 3.9 ± 1.9 ) × 10 <sup>-3</sup>	
$\Gamma_{48}$ $D^-\pi^+\rho^0$	( 1.1 ± 1.0 ) × 10 <sup>-3</sup>	
$\Gamma_{49}$ $D^-\bar{a}_1(1260)^+$	( 6.0 ± 3.3 ) × 10 <sup>-3</sup>	
$\Gamma_{50}$ $D^*(2010)^-\pi^+\pi^0$	( 1.5 ± 0.5 ) %	
$\Gamma_{51}$ $D^*(2010)^-\rho^+$	( 2.2 ± 1.8 ) × 10 <sup>-3</sup>	S=5.2
$\Gamma_{52}$ $D^*(2010)^-K^+$	( 2.12 ± 0.15 ) × 10 <sup>-4</sup>	
$\Gamma_{53}$ $D^*(2010)^-K^0\pi^+$	( 3.0 ± 0.8 ) × 10 <sup>-4</sup>	
$\Gamma_{54}$ $D^*(2010)^-K^*(892)^+$	( 3.3 ± 0.6 ) × 10 <sup>-4</sup>	
$\Gamma_{55}$ $D^*(2010)^-K^+\bar{K}^0$	< 4.7	× 10 <sup>-4</sup> CL=90%
$\Gamma_{56}$ $D^*(2010)^-K^+\bar{K}^*(892)^0$	( 1.29 ± 0.33 ) × 10 <sup>-3</sup>	
$\Gamma_{57}$ $D^*(2010)^-\pi^+\pi^+\pi^-$	( 7.0 ± 0.8 ) × 10 <sup>-3</sup>	S=1.3
$\Gamma_{58}$ $(D^*(2010)^-\pi^+\pi^+\pi^-)$ non-resonant	( 0.0 ± 2.5 ) × 10 <sup>-3</sup>	
$\Gamma_{59}$ $D^*(2010)^-\pi^+\rho^0$	( 5.7 ± 3.2 ) × 10 <sup>-3</sup>	
$\Gamma_{60}$ $D^*(2010)^-\bar{a}_1(1260)^+$	( 1.30 ± 0.27 ) %	
$\Gamma_{61}$ $\bar{D}_1(2420)^0\pi^-\pi^+$ , $\bar{D}_1^0 \rightarrow D^{*-}\pi^+$	( 1.4 ± 0.4 ) × 10 <sup>-4</sup>	
$\Gamma_{62}$ $D^*(2010)^-K^+\pi^-\pi^+$	( 4.5 ± 0.7 ) × 10 <sup>-4</sup>	
$\Gamma_{63}$ $D^*(2010)^-\pi^+\pi^+\pi^-\pi^0$	( 1.76 ± 0.27 ) %	
$\Gamma_{64}$ $D^{*-}3\pi^+2\pi^-$	( 4.7 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{65}$ $\bar{D}^*(2010)^-\omega\pi^+$	( 2.46 ± 0.18 ) × 10 <sup>-3</sup>	S=1.2
$\Gamma_{66}$ $D_1(2430)^0\omega$ , $D_1^0 \rightarrow D^{*-}\pi^+$	( 2.7 ± 0.8 ) × 10 <sup>-4</sup>	
$\Gamma_{67}$ $\bar{D}^{*-}\rho(1450)^+$	( 1.07 ± 0.40 ) × 10 <sup>-3</sup>	
$\Gamma_{68}$ $\bar{D}_1(2420)^0\omega$	( 7.0 ± 2.2 ) × 10 <sup>-5</sup>	
$\Gamma_{69}$ $\bar{D}_2^*(2460)^0\omega$	( 4.0 ± 1.4 ) × 10 <sup>-5</sup>	
$\Gamma_{70}$ $\bar{D}^{*-}b_1(1235)^-$ , $b_1^- \rightarrow \omega\pi^-$	< 7	× 10 <sup>-5</sup> CL=90%
$\Gamma_{71}$ $\bar{D}^{*-}\pi^+$	[b] ( 1.9 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{72}$ $D_1(2420)^-\pi^+$ , $D_1^- \rightarrow D^{*-}\pi^+$	( 9.9 ± 2.0 ) × 10 <sup>-5</sup>	
$\Gamma_{73}$ $D_1(2420)^-\pi^+$ , $D_1^- \rightarrow D^{*-}\pi^+\pi^-$	< 3.3	× 10 <sup>-5</sup> CL=90%
$\Gamma_{74}$ $\bar{D}_2^*(2460)^-\pi^+$ , $(D_2^*)^- \rightarrow D^{*-}\pi^+\pi^-$	( 2.38 ± 0.16 ) × 10 <sup>-4</sup>	
$\Gamma_{75}$ $\bar{D}_0^*(2400)^-\pi^+$ , $(D_0^*)^- \rightarrow D^{*-}\pi^+\pi^-$	( 7.6 ± 0.8 ) × 10 <sup>-5</sup>	
$\Gamma_{76}$ $D_2^*(2460)^-\pi^+$ , $(D_2^*)^- \rightarrow D^{*-}\pi^+\pi^-$	< 2.4	× 10 <sup>-5</sup> CL=90%
$\Gamma_{77}$ $\bar{D}_2^*(2460)^-\rho^+$	< 4.9	× 10 <sup>-3</sup> CL=90%



$\Gamma_{78}$	$D^0 \bar{D}^0$	$(1.4 \pm 0.7) \times 10^{-5}$				$\Gamma_{128}$	$D_s^{*-} \pi^+ K^0$	$< 1.10$	$\times 10^{-4}$	CL=90%
$\Gamma_{79}$	$D^{*0} \bar{D}^0$	$< 2.9$	$\times 10^{-4}$	CL=90%		$\Gamma_{129}$	$D_s^- K^+ \pi^+ \pi^-$	$(1.7 \pm 0.5) \times 10^{-4}$		
$\Gamma_{80}$	$D^- D^+$	$(2.11 \pm 0.18) \times 10^{-4}$				$\Gamma_{130}$	$D_s^- \pi^+ K^*(892)^0$	$< 3.0$	$\times 10^{-3}$	CL=90%
$\Gamma_{81}$	$D^\pm D^{*\mp} (CP\text{-averaged})$	$(6.1 \pm 0.6) \times 10^{-4}$				$\Gamma_{131}$	$D_s^{*-} \pi^+ K^*(892)^0$	$< 1.6$	$\times 10^{-3}$	CL=90%
$\Gamma_{82}$	$D^- D_s^+$	$(7.2 \pm 0.8) \times 10^{-3}$				$\Gamma_{132}$	$\bar{D}^0 K^0$	$(5.2 \pm 0.7) \times 10^{-5}$		
$\Gamma_{83}$	$D^*(2010)^- D_s^+$	$(8.0 \pm 1.1) \times 10^{-3}$				$\Gamma_{133}$	$\bar{D}^0 K^+ \pi^-$	$(8.8 \pm 1.7) \times 10^{-5}$		
$\Gamma_{84}$	$D^- D_s^{*+}$	$(7.4 \pm 1.6) \times 10^{-3}$				$\Gamma_{134}$	$\bar{D}^0 K^*(892)^0$	$(4.5 \pm 0.6) \times 10^{-5}$		
$\Gamma_{85}$	$D^*(2010)^- D_s^{*+}$	$(1.77 \pm 0.14) \%$				$\Gamma_{135}$	$\bar{D}^0 K^*(1410)^0$	$< 6.7$	$\times 10^{-5}$	CL=90%
$\Gamma_{86}$	$D_{s0}(2317)^- K^+, D_{s0}^- \rightarrow D_s^- \pi^0$	$(4.2 \pm 1.4) \times 10^{-5}$				$\Gamma_{136}$	$\bar{D}^0 K_0^*(1430)^0$	$(7 \pm 7) \times 10^{-6}$		
$\Gamma_{87}$	$D_{s0}(2317)^- \pi^+, D_{s0}^- \rightarrow D_s^- \pi^0$	$< 2.5$	$\times 10^{-5}$	CL=90%		$\Gamma_{137}$	$\bar{D}^0 K_2^*(1430)^0$	$(2.1 \pm 0.9) \times 10^{-5}$		
$\Gamma_{88}$	$D_{sJ}(2457)^- K^+, D_{sJ}^- \rightarrow D_s^- \pi^0$	$< 9.4$	$\times 10^{-6}$	CL=90%		$\Gamma_{138}$	$D_0^*(2400)^-, D_0^{*-} \rightarrow \bar{D}^0 \pi^-$	$(1.9 \pm 0.9) \times 10^{-5}$		
$\Gamma_{89}$	$D_{sJ}(2457)^- \pi^+, D_{sJ}^- \rightarrow D_s^- \pi^0$	$< 4.0$	$\times 10^{-6}$	CL=90%		$\Gamma_{139}$	$D_2^*(2460)^- K^+, D_2^{*-} \rightarrow \bar{D}^0 \pi^-$	$(2.03 \pm 0.35) \times 10^{-5}$		
$\Gamma_{90}$	$D_s^- D_s^+$	$< 3.6$	$\times 10^{-5}$	CL=90%		$\Gamma_{140}$	$D_3^*(2760)^- K^+, D_3^{*-} \rightarrow \bar{D}^0 \pi^-$	$< 1.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{91}$	$D_s^{*-} D_s^+$	$< 1.3$	$\times 10^{-4}$	CL=90%		$\Gamma_{141}$	$\bar{D}^0 K^+ \pi^- \text{ non-resonant}$	$< 3.7$	$\times 10^{-5}$	CL=90%
$\Gamma_{92}$	$D_s^{*-} D_s^{*+}$	$< 2.4$	$\times 10^{-4}$	CL=90%		$\Gamma_{142}$	$[K^+ K^-]_D K^*(892)^0$	$(4.7 \pm 0.9) \times 10^{-5}$		
$\Gamma_{93}$	$D_{s0}^*(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0$	$(1.04 \pm 0.17) \times 10^{-3}$	S=1.1			$\Gamma_{143}$	$[\pi^+ \pi^-]_D K^*(892)^0$	$(5.5 \pm 1.4) \times 10^{-5}$		
$\Gamma_{94}$	$D_{s0}(2317)^+ D^-, D_{s0}^+ \rightarrow D_s^{*+} \gamma$	$< 9.5$	$\times 10^{-4}$	CL=90%		$\Gamma_{144}$	$\bar{D}^0 \rho^0$	$(2.63 \pm 0.14) \times 10^{-4}$		
$\Gamma_{95}$	$D_{s0}(2317)^+ D^*(2010)^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0$	$(1.5 \pm 0.6) \times 10^{-3}$				$\Gamma_{145}$	$\bar{D}^0 \rho^0$	$(3.21 \pm 0.21) \times 10^{-4}$		
$\Gamma_{96}$	$D_{sJ}(2457)^+ D^-$	$(3.5 \pm 1.1) \times 10^{-3}$				$\Gamma_{146}$	$\bar{D}^0 f_2$	$(1.56 \pm 0.21) \times 10^{-4}$		
$\Gamma_{97}$	$D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \gamma$	$(6.5 \pm 1.7) \times 10^{-4}$				$\Gamma_{147}$	$\bar{D}^0 \eta$	$(2.36 \pm 0.32) \times 10^{-4}$	S=2.5	
$\Gamma_{98}$	$D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^{*+} \gamma$	$< 6.0$	$\times 10^{-4}$	CL=90%		$\Gamma_{148}$	$\bar{D}^0 \eta'$	$(1.38 \pm 0.16) \times 10^{-4}$	S=1.3	
$\Gamma_{99}$	$D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \pi^+ \pi^-$	$< 2.0$	$\times 10^{-4}$	CL=90%		$\Gamma_{149}$	$\bar{D}^0 \omega$	$(2.54 \pm 0.16) \times 10^{-4}$		
$\Gamma_{100}$	$D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \pi^0$	$< 3.6$	$\times 10^{-4}$	CL=90%		$\Gamma_{150}$	$D^0 \phi$	$< 1.16$	$\times 10^{-5}$	CL=90%
$\Gamma_{101}$	$D^*(2010)^- D_{sJ}(2457)^+$	$(9.3 \pm 2.2) \times 10^{-3}$				$\Gamma_{151}$	$D^0 K^+ \pi^-$	$(5.3 \pm 3.2) \times 10^{-6}$		
$\Gamma_{102}$	$D_{sJ}(2457)^+ D^*(2010), D_{sJ}^+ \rightarrow D_s^+ \gamma$	$(2.3 \pm 0.9) \times 10^{-3}$				$\Gamma_{152}$	$D^0 K^*(892)^0$	$< 1.1$	$\times 10^{-5}$	CL=90%
$\Gamma_{103}$	$D^- D_{s1}(2536)^+, D_{s1}^- \rightarrow D^{*0} K^+ + D^{*+} K^0$	$(2.8 \pm 0.7) \times 10^{-4}$				$\Gamma_{153}$	$\bar{D}^{*0} \gamma$	$< 2.5$	$\times 10^{-5}$	CL=90%
$\Gamma_{104}$	$D^- D_{s1}(2536)^+, D_{s1}^- \rightarrow D^{*0} K^+$	$(1.7 \pm 0.6) \times 10^{-4}$				$\Gamma_{154}$	$\bar{D}^*(2007)^0 \pi^0$	$(2.2 \pm 0.6) \times 10^{-4}$	S=2.6	
$\Gamma_{105}$	$D^- D_{s1}(2536)^+, D_{s1}^- \rightarrow D^{*+} K^0$	$(2.6 \pm 1.1) \times 10^{-4}$				$\Gamma_{155}$	$\bar{D}^*(2007)^0 \rho^0$	$< 5.1$	$\times 10^{-4}$	CL=90%
$\Gamma_{106}$	$D^*(2010)^- D_{s1}(2536)^+, D_{s1}^- \rightarrow D^{*0} K^+ + D^{*+} K^0$	$(5.0 \pm 1.4) \times 10^{-4}$				$\Gamma_{156}$	$\bar{D}^*(2007)^0 \eta$	$(2.3 \pm 0.6) \times 10^{-4}$	S=2.8	
$\Gamma_{107}$	$D^*(2010)^- D_{s1}(2536)^+, D_{s1}^- \rightarrow D^{*0} K^+$	$(3.3 \pm 1.1) \times 10^{-4}$				$\Gamma_{157}$	$\bar{D}^*(2007)^0 \eta'$	$(1.40 \pm 0.22) \times 10^{-4}$		
$\Gamma_{108}$	$D^{*-} D_{s1}(2536)^+, D_{s1}^- \rightarrow D^{*+} K^0$	$(5.0 \pm 1.7) \times 10^{-4}$				$\Gamma_{158}$	$\bar{D}^*(2007)^0 \pi^+ \pi^-$	$(6.2 \pm 2.2) \times 10^{-4}$		
$\Gamma_{109}$	$D^- D_{sJ}(2573)^+, D_{sJ}^- \rightarrow D^0 K^+$	$(3.4 \pm 1.8) \times 10^{-5}$				$\Gamma_{159}$	$\bar{D}^*(2007)^0 K^0$	$(3.6 \pm 1.2) \times 10^{-5}$		
$\Gamma_{110}$	$D^*(2010)^- D_{sJ}(2573)^+, D_{sJ}^- \rightarrow D^0 K^+$	$< 2$	$\times 10^{-4}$	CL=90%		$\Gamma_{160}$	$\bar{D}^*(2007)^0 K^*(892)^0$	$< 6.9$	$\times 10^{-5}$	CL=90%
$\Gamma_{111}$	$D^- D_{sJ}(2700)^+, D_{sJ}^- \rightarrow D^0 K^+$	$(7.1 \pm 1.2) \times 10^{-4}$				$\Gamma_{161}$	$D^*(2007)^0 K^*(892)^0$	$< 4.0$	$\times 10^{-5}$	CL=90%
$\Gamma_{112}$	$D^+ \pi^-$	$(7.4 \pm 1.3) \times 10^{-7}$				$\Gamma_{162}$	$D^*(2007)^0 \pi^+ \pi^+ \pi^- \pi^-$	$(2.7 \pm 0.5) \times 10^{-3}$		
$\Gamma_{113}$	$D_s^+ \pi^-$	$(2.16 \pm 0.26) \times 10^{-5}$				$\Gamma_{163}$	$D^*(2010)^+ D^*(2010)^-$	$(8.0 \pm 0.6) \times 10^{-4}$		
$\Gamma_{114}$	$D_s^{*+} \pi^-$	$(2.1 \pm 0.4) \times 10^{-5}$	S=1.4			$\Gamma_{164}$	$\bar{D}^*(2007)^0 \omega$	$(3.6 \pm 1.1) \times 10^{-4}$	S=3.1	
$\Gamma_{115}$	$D_s^+ \rho^-$	$< 2.4$	$\times 10^{-5}$	CL=90%		$\Gamma_{165}$	$D^*(2010)^+ D^-$	$(6.1 \pm 1.5) \times 10^{-4}$	S=1.6	
$\Gamma_{116}$	$D_s^{*+} \rho^-$	$(4.1 \pm 1.3) \times 10^{-5}$				$\Gamma_{166}$	$D^*(2007)^0 \bar{D}^*(2007)^0$	$< 9$	$\times 10^{-5}$	CL=90%
$\Gamma_{117}$	$D_s^+ a_0^-$	$< 1.9$	$\times 10^{-5}$	CL=90%		$\Gamma_{167}$	$D^- D^0 K^+$	$(1.07 \pm 0.11) \times 10^{-3}$		
$\Gamma_{118}$	$D_s^{*+} a_0^-$	$< 3.6$	$\times 10^{-5}$	CL=90%		$\Gamma_{168}$	$D^- D^*(2007)^0 K^+$	$(3.5 \pm 0.4) \times 10^{-3}$		
$\Gamma_{119}$	$D_s^+ a_1(1260)^-$	$< 2.1$	$\times 10^{-3}$	CL=90%		$\Gamma_{169}$	$D^*(2010)^- D^0 K^+$	$(2.47 \pm 0.21) \times 10^{-3}$		
$\Gamma_{120}$	$D_s^{*+} a_1(1260)^-$	$< 1.7$	$\times 10^{-3}$	CL=90%		$\Gamma_{170}$	$D^*(2010)^- D^*(2007)^0 K^+$	$(1.06 \pm 0.09) \%$		
$\Gamma_{121}$	$D_s^+ a_2^-$	$< 1.9$	$\times 10^{-4}$	CL=90%		$\Gamma_{171}$	$D^- D^+ K^0$	$(7.5 \pm 1.7) \times 10^{-4}$		
$\Gamma_{122}$	$D_s^{*+} a_2^-$	$< 2.0$	$\times 10^{-4}$	CL=90%		$\Gamma_{172}$	$D^*(2010)^- D^+ K^0 + D^- D^*(2010)^+ K^0$	$(6.4 \pm 0.5) \times 10^{-3}$		
$\Gamma_{123}$	$D_s^- K^+$	$(2.7 \pm 0.5) \times 10^{-5}$	S=2.7			$\Gamma_{173}$	$D^*(2010)^- D^*(2010)^+ K^0$	$(8.1 \pm 0.7) \times 10^{-3}$		
$\Gamma_{124}$	$D_s^{*-} K^+$	$(2.19 \pm 0.30) \times 10^{-5}$				$\Gamma_{174}$	$D^{*-} D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0$	$(8.0 \pm 2.4) \times 10^{-4}$		
$\Gamma_{125}$	$D_s^- K^*(892)^+$	$(3.5 \pm 1.0) \times 10^{-5}$				$\Gamma_{175}$	$\bar{D}^0 D^0 K^0$	$(2.7 \pm 1.1) \times 10^{-4}$		
$\Gamma_{126}$	$D_s^{*-} K^*(892)^+$	$(3.2 \pm 1.5) \times 10^{-5}$				$\Gamma_{176}$	$\bar{D}^0 D^*(2007)^0 K^0 + \bar{D}^*(2007)^0 D^0 K^0$	$(1.1 \pm 0.5) \times 10^{-3}$		
$\Gamma_{127}$	$D_s^- \pi^+ K^0$	$(9.7 \pm 1.4) \times 10^{-5}$				$\Gamma_{177}$	$\bar{D}^*(2007)^0 D^*(2007)^0 K^0$	$(2.4 \pm 0.9) \times 10^{-3}$		
						$\Gamma_{178}$	$(\bar{D} + \bar{D}^*)(D + D^*)K$	$(3.68 \pm 0.26) \%$		

## Charmonium modes

$\Gamma_{179}$	$\eta_c K^0$	$(8.0 \pm 1.2) \times 10^{-4}$		
$\Gamma_{180}$	$\eta_c K^*(892)^0$	$(6.3 \pm 0.9) \times 10^{-4}$		
$\Gamma_{181}$	$\eta_c(2S) K^{*0}$	$< 3.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{182}$	$h_c(1P) K^{*0}$	$< 4$	$\times 10^{-4}$	CL=90%
$\Gamma_{183}$	$J/\psi(1S) K^0$	$(8.73 \pm 0.32) \times 10^{-4}$		
$\Gamma_{184}$	$J/\psi(1S) K^+ \pi^-$	$(1.15 \pm 0.05) \times 10^{-3}$		
$\Gamma_{185}$	$J/\psi(1S) K^*(892)^0$	$(1.28 \pm 0.05) \times 10^{-3}$		
$\Gamma_{186}$	$J/\psi(1S) \eta K_S^0$	$(5.4 \pm 0.9) \times 10^{-5}$		
$\Gamma_{187}$	$J/\psi(1S) \eta' K_S^0$	$< 2.5$	$\times 10^{-5}$	CL=90%
$\Gamma_{188}$	$J/\psi(1S) \phi K^0$	$(4.9 \pm 1.0) \times 10^{-5}$	S=1.3	
$\Gamma_{189}$	$J/\psi(1S) \omega K^0$	$(2.3 \pm 0.4) \times 10^{-4}$		
$\Gamma_{190}$	$X(3872) K^0, X \rightarrow J/\psi \omega$	$(6.0 \pm 3.2) \times 10^{-6}$		
$\Gamma_{191}$	$X(3915), X \rightarrow J/\psi \omega$	$(2.1 \pm 0.9) \times 10^{-5}$		
$\Gamma_{192}$	$J/\psi(1S) K(1270)^0$	$(1.3 \pm 0.5) \times 10^{-3}$		
$\Gamma_{193}$	$J/\psi(1S) \pi^0$	$(1.76 \pm 0.16) \times 10^{-5}$	S=1.1	
$\Gamma_{194}$	$J/\psi(1S) \eta$	$(1.08 \pm 0.24) \times 10^{-5}$	S=1.5	
$\Gamma_{195}$	$J/\psi(1S) \pi^+ \pi^-$	$(4.03 \pm 0.18) \times 10^{-5}$		

## Meson Particle Listings

 $B^0$ 

$\Gamma_{196}$	$J/\psi(1S)\pi^+\pi^-$ nonresonant	$< 1.2 \times 10^{-5}$	CL=90%	$\Gamma_{257}$	$\eta' K_S^*(1430)^0$	$(1.37 \pm 0.32) \times 10^{-5}$	
$\Gamma_{197}$	$J/\psi(1S)f_0(500), f_0 \rightarrow \pi\pi$	$(8.1 \pm_{-0.9}^{+1.1}) \times 10^{-6}$		$\Gamma_{258}$	$\eta K^0$	$(1.23 \pm_{-0.24}^{+0.27}) \times 10^{-6}$	
$\Gamma_{198}$	$J/\psi(1S)f_2$	$(3.3 \pm_{-0.6}^{+0.5}) \times 10^{-6}$	S=1.6	$\Gamma_{259}$	$\eta K^*(892)^0$	$(1.59 \pm 0.10) \times 10^{-5}$	
$\Gamma_{199}$	$J/\psi(1S)\rho^0$	$(2.54 \pm 0.14) \times 10^{-5}$		$\Gamma_{260}$	$\eta K_0^*(1430)^0$	$(1.10 \pm 0.22) \times 10^{-5}$	
$\Gamma_{200}$	$J/\psi(1S)f_0(980), f_0 \rightarrow \pi^+\pi^-$	$< 1.1 \times 10^{-6}$	CL=90%	$\Gamma_{261}$	$\eta K_S^*(1430)^0$	$(9.6 \pm 2.1) \times 10^{-6}$	
$\Gamma_{201}$	$J/\psi(1S)\rho(1450)^0, \rho^0 \rightarrow \pi\pi$	$(3.0 \pm_{-0.7}^{+1.6}) \times 10^{-6}$		$\Gamma_{262}$	$\omega K^0$	$(4.8 \pm 0.4) \times 10^{-6}$	
$\Gamma_{202}$	$J/\psi\rho(1700)^0, \rho^0 \rightarrow \pi^+\pi^-$	$(2.0 \pm 1.3) \times 10^{-6}$		$\Gamma_{263}$	$a_0(980)^0 K^0, a_0^0 \rightarrow \eta\pi^0$	$< 7.8 \times 10^{-6}$	CL=90%
$\Gamma_{203}$	$J/\psi(1S)\omega$	$(1.8 \pm_{-0.5}^{+0.7}) \times 10^{-5}$		$\Gamma_{264}$	$b_1^0 K^0, b_1^0 \rightarrow \omega\pi^0$	$< 7.8 \times 10^{-6}$	CL=90%
$\Gamma_{204}$	$J/\psi(1S)K^+K^-$	$(2.6 \pm 0.4) \times 10^{-6}$		$\Gamma_{265}$	$a_0(980)^\pm K^\mp, a_0^\pm \rightarrow \eta\pi^\pm$	$< 1.9 \times 10^{-6}$	CL=90%
$\Gamma_{205}$	$J/\psi(1S)a_0(980), a_0 \rightarrow K^+K^-$	$(4.7 \pm 3.4) \times 10^{-7}$		$\Gamma_{266}$	$b_1^- K^+, b_1^- \rightarrow \omega\pi^-$	$(7.4 \pm 1.4) \times 10^{-6}$	
$\Gamma_{206}$	$J/\psi(1S)\phi$	$< 1.9 \times 10^{-7}$	CL=90%	$\Gamma_{267}$	$b_1^0 K^{*0}, b_1^0 \rightarrow \omega\pi^0$	$< 8.0 \times 10^{-6}$	CL=90%
$\Gamma_{207}$	$J/\psi(1S)\eta'(958)$	$(7.6 \pm 2.4) \times 10^{-6}$		$\Gamma_{268}$	$b_1^- K^{*+}, b_1^- \rightarrow \omega\pi^-$	$< 5.0 \times 10^{-6}$	CL=90%
$\Gamma_{208}$	$J/\psi(1S)K^0\pi^+\pi^-$	$(4.4 \pm 0.4) \times 10^{-4}$		$\Gamma_{269}$	$a_0(1450)^\pm K^\mp, a_0^\pm \rightarrow \eta\pi^\pm$	$< 3.1 \times 10^{-6}$	CL=90%
$\Gamma_{209}$	$J/\psi(1S)K^0K^-\pi^+ + \text{c.c.}$	$< 2.1 \times 10^{-5}$	CL=90%	$\Gamma_{270}$	$K_S^0 X^0$ (Familon)	$< 5.3 \times 10^{-5}$	CL=90%
$\Gamma_{210}$	$J/\psi(1S)K^0K^+K^-$	$(2.5 \pm 0.7) \times 10^{-5}$	S=1.8	$\Gamma_{271}$	$\omega K^*(892)^0$	$(2.0 \pm 0.5) \times 10^{-6}$	
$\Gamma_{211}$	$J/\psi(1S)K^0K^\pm\pi^\mp$			$\Gamma_{272}$	$\omega(K\pi)_0^{*0}$	$(1.84 \pm 0.25) \times 10^{-5}$	
$\Gamma_{212}$	$J/\psi(1S)K^0\rho^0$	$(5.4 \pm 3.0) \times 10^{-4}$		$\Gamma_{273}$	$\omega K_0^*(1430)^0$	$(1.60 \pm 0.34) \times 10^{-5}$	
$\Gamma_{213}$	$J/\psi(1S)K^*(892)^\pm\pi^\mp$	$(8 \pm 4) \times 10^{-4}$		$\Gamma_{274}$	$\omega K_S^*(1430)^0$	$(1.01 \pm 0.23) \times 10^{-5}$	
$\Gamma_{214}$	$J/\psi(1S)\pi^+\pi^-\pi^+\pi^-$	$(1.45 \pm 0.13) \times 10^{-5}$		$\Gamma_{275}$	$\omega K^+\pi^-$ nonresonant	$(5.1 \pm 1.0) \times 10^{-6}$	
$\Gamma_{215}$	$J/\psi(1S)f_1(1285)$	$(8.4 \pm 2.1) \times 10^{-6}$		$\Gamma_{276}$	$K^+\pi^-\pi^0$	$(3.78 \pm 0.32) \times 10^{-5}$	
$\Gamma_{216}$	$J/\psi(1S)K^*(892)^0\pi^+\pi^-$	$(6.6 \pm 2.2) \times 10^{-4}$		$\Gamma_{277}$	$K^+\rho^-$	$(7.0 \pm 0.9) \times 10^{-6}$	
$\Gamma_{217}$	$X(3872)^- K^+$	$< 5 \times 10^{-4}$	CL=90%	$\Gamma_{278}$	$K^+\rho(1450)^-$	$(2.4 \pm 1.2) \times 10^{-6}$	
$\Gamma_{218}$	$X(3872)^- K^+, X(3872)^- \rightarrow J/\psi(1S)\pi^-\pi^0$	$[c] < 4.2 \times 10^{-6}$	CL=90%	$\Gamma_{279}$	$K^+\rho(1700)^-$	$(6 \pm 7) \times 10^{-7}$	
$\Gamma_{219}$	$X(3872)K^0, X \rightarrow J/\psi\pi^+\pi^-$	$(4.3 \pm 1.3) \times 10^{-6}$		$\Gamma_{280}$	$(K^+\pi^-\pi^0)$ non-resonant	$(2.8 \pm 0.6) \times 10^{-6}$	
$\Gamma_{220}$	$X(3872)K^0, X \rightarrow J/\psi\gamma$	$< 2.4 \times 10^{-6}$	CL=90%	$\Gamma_{281}$	$(K\pi)_0^{*+}\pi^-, (K\pi)_0^{*+} \rightarrow K^+\pi^0$	$(8.6 \pm 1.7) \times 10^{-6}$	
$\Gamma_{221}$	$X(3872)K^*(892)^0, X \rightarrow J/\psi\gamma$	$< 2.8 \times 10^{-6}$	CL=90%	$\Gamma_{282}$	$K^+\pi^-\pi^0$		
$\Gamma_{222}$	$X(3872)K^0, X \rightarrow \psi(2S)\gamma$	$< 6.62 \times 10^{-6}$	CL=90%	$\Gamma_{283}$	$K_S^*(1430)^0\pi^0$	$< 4.0 \times 10^{-6}$	CL=90%
$\Gamma_{223}$	$X(3872)K^*(892)^0, X \rightarrow \psi(2S)\gamma$	$< 4.4 \times 10^{-6}$	CL=90%	$\Gamma_{284}$	$K^*(1680)^0\pi^0$	$< 7.5 \times 10^{-6}$	CL=90%
$\Gamma_{224}$	$X(3872)K^0, X \rightarrow D^0\bar{D}^0\pi^0$	$(1.7 \pm 0.8) \times 10^{-4}$		$\Gamma_{285}$	$K^{*0}\pi^0$	$[d] (6.1 \pm 1.6) \times 10^{-6}$	
$\Gamma_{225}$	$X(3872)K^0, X \rightarrow \bar{D}^{*0}D^0$	$(1.2 \pm 0.4) \times 10^{-4}$		$\Gamma_{286}$	$K^0\pi^+\pi^-$	$(5.20 \pm 0.24) \times 10^{-5}$	S=1.3
$\Gamma_{226}$	$X(3872)K^+\pi^-, X \rightarrow J/\psi\pi^+\pi^-$	$(7.9 \pm 1.4) \times 10^{-6}$		$\Gamma_{287}$	$K^0\pi^+\pi^-$ non-resonant	$(1.47 \pm_{-0.26}^{+0.40}) \times 10^{-5}$	S=2.1
$\Gamma_{227}$	$X(3872)K^*(982)^0, X \rightarrow J/\psi\pi^+\pi^-$	$(4.0 \pm 1.5) \times 10^{-6}$		$\Gamma_{288}$	$K^0\rho^0$	$(4.7 \pm 0.6) \times 10^{-6}$	
$\Gamma_{228}$	$X(4430)^\pm K^\mp, X^\pm \rightarrow \psi(2S)\pi^\pm$	$(6.0 \pm_{-2.4}^{+3.0}) \times 10^{-5}$		$\Gamma_{289}$	$K^*(892)^\pm\pi^\mp$	$(8.4 \pm 0.8) \times 10^{-6}$	
$\Gamma_{229}$	$X(4430)^\pm K^\mp, X^\pm \rightarrow J/\psi\pi^\pm$	$(5.4 \pm_{-1.2}^{+4.0}) \times 10^{-6}$		$\Gamma_{290}$	$K_0^*(1430)^\pm\pi^\mp$	$(3.3 \pm 0.7) \times 10^{-5}$	S=2.0
$\Gamma_{230}$	$X(3900)^\pm K^\mp, X^\pm \rightarrow J/\psi\pi^\pm$	$< 9 \times 10^{-7}$		$\Gamma_{291}$	$K^{*+}\pi^-$	$[d] (5.1 \pm 1.6) \times 10^{-6}$	
$\Gamma_{231}$	$X(4200)^\pm K^\mp, X^\pm \rightarrow J/\psi\pi^\pm$	$(2.2 \pm_{-0.8}^{+1.3}) \times 10^{-5}$		$\Gamma_{292}$	$K^*(1410)^\pm\pi^\mp, K^{*+} \rightarrow K^0\pi^+$	$< 3.8 \times 10^{-6}$	CL=90%
$\Gamma_{232}$	$J/\psi(1S)p\bar{p}$	$< 5.2 \times 10^{-7}$	CL=90%	$\Gamma_{293}$	$f_0(980)K^0, f_0 \rightarrow \pi^+\pi^-$	$(7.0 \pm 0.9) \times 10^{-6}$	
$\Gamma_{233}$	$J/\psi(1S)\gamma$	$< 1.5 \times 10^{-6}$	CL=90%	$\Gamma_{294}$	$f_2(1270)K^0$	$(2.7 \pm_{-1.2}^{+1.3}) \times 10^{-6}$	
$\Gamma_{234}$	$J/\psi(1S)\bar{D}^0$	$< 1.3 \times 10^{-5}$	CL=90%	$\Gamma_{295}$	$f_x(1300)K^0, f_x \rightarrow \pi^+\pi^-$	$(1.8 \pm 0.7) \times 10^{-6}$	
$\Gamma_{235}$	$\psi(2S)\pi^0$	$(1.17 \pm 0.19) \times 10^{-5}$		$\Gamma_{296}$	$K^*(892)^0\pi^0$	$(3.3 \pm 0.6) \times 10^{-6}$	
$\Gamma_{236}$	$\psi(2S)K^0$	$(5.8 \pm 0.5) \times 10^{-4}$		$\Gamma_{297}$	$K_S^*(1430)^\pm\pi^\mp$	$< 6 \times 10^{-6}$	CL=90%
$\Gamma_{237}$	$\psi(3770)K^0, \psi \rightarrow \bar{D}^0D^0$	$< 1.23 \times 10^{-4}$	CL=90%	$\Gamma_{298}$	$K^*(1680)^\pm\pi^\mp$	$< 1.0 \times 10^{-5}$	CL=90%
$\Gamma_{238}$	$\psi(3770)K^0, \psi \rightarrow D^-D^+$	$< 1.88 \times 10^{-4}$	CL=90%	$\Gamma_{299}$	$K^+\pi^-\pi^+\pi^-$	$[e] < 2.3 \times 10^{-4}$	CL=90%
$\Gamma_{239}$	$\psi(2S)\pi^+\pi^-$	$(2.3 \pm 0.4) \times 10^{-5}$		$\Gamma_{300}$	$\rho^0 K^+\pi^-$	$(2.8 \pm 0.7) \times 10^{-6}$	
$\Gamma_{240}$	$\psi(2S)K^+\pi^-$	$(5.8 \pm 0.4) \times 10^{-4}$		$\Gamma_{301}$	$f_0(980)K^+\pi^-, f_0 \rightarrow \pi\pi$	$(1.4 \pm_{-0.6}^{+0.5}) \times 10^{-6}$	
$\Gamma_{241}$	$\psi(2S)K^*(892)^0$	$(5.9 \pm 0.4) \times 10^{-4}$		$\Gamma_{302}$	$K^+\pi^-\pi^+\pi^-$ nonresonant	$< 2.1 \times 10^{-6}$	CL=90%
$\Gamma_{242}$	$\chi_{c0}K^0$	$(1.47 \pm 0.27) \times 10^{-4}$		$\Gamma_{303}$	$K^*(892)^0\pi^+\pi^-$	$(5.5 \pm 0.5) \times 10^{-5}$	
$\Gamma_{243}$	$\chi_{c0}K^*(892)^0$	$(1.7 \pm 0.4) \times 10^{-4}$		$\Gamma_{304}$	$K^*(892)^0\rho^0$	$(3.9 \pm 1.3) \times 10^{-6}$	S=1.9
$\Gamma_{244}$	$\chi_{c2}K^0$	$< 1.5 \times 10^{-5}$	CL=90%	$\Gamma_{305}$	$K^*(892)^0 f_0(980), f_0 \rightarrow \pi\pi$	$(3.9 \pm_{-1.8}^{+2.1}) \times 10^{-6}$	S=3.9
$\Gamma_{245}$	$\chi_{c2}K^*(892)^0$	$(4.9 \pm 1.2) \times 10^{-5}$	S=1.1	$\Gamma_{306}$	$K_1(1270)^\pm\pi^\mp$	$< 3.0 \times 10^{-5}$	CL=90%
$\Gamma_{246}$	$\chi_{c1}\pi^0$	$(1.12 \pm 0.28) \times 10^{-5}$		$\Gamma_{307}$	$K_1(1400)^\pm\pi^\mp$	$< 2.7 \times 10^{-5}$	CL=90%
$\Gamma_{247}$	$\chi_{c1}K^0$	$(3.93 \pm 0.27) \times 10^{-4}$		$\Gamma_{308}$	$a_1(1260)^- K^+$	$[e] (1.6 \pm 0.4) \times 10^{-5}$	
$\Gamma_{248}$	$\chi_{c1}K^-\pi^+$	$(3.8 \pm 0.4) \times 10^{-4}$		$\Gamma_{309}$	$K^*(892)^\pm\rho^-$	$(1.03 \pm 0.26) \times 10^{-5}$	
$\Gamma_{249}$	$\chi_{c1}K^*(892)^0$	$(2.39 \pm 0.19) \times 10^{-4}$	S=1.2	$\Gamma_{310}$	$K_0^*(1430)^\pm\rho^-$	$(2.8 \pm 1.2) \times 10^{-5}$	
$\Gamma_{250}$	$X(4051)^+ K^-, X^+ \rightarrow \chi_{c1}\pi^+$	$(3.0 \pm_{-1.8}^{+4.0}) \times 10^{-5}$		$\Gamma_{311}$	$K_1(1400)^0\rho^0$	$< 3.0 \times 10^{-3}$	CL=90%
$\Gamma_{251}$	$X(4248)^+ K^-, X^+ \rightarrow \chi_{c1}\pi^+$	$(4.0 \pm_{-1.0}^{+20.0}) \times 10^{-5}$		$\Gamma_{312}$	$K_0^*(1430)^0\rho^0$	$(2.7 \pm 0.6) \times 10^{-5}$	
$\Gamma_{252}$	$K^+\pi^-$	$(1.96 \pm 0.05) \times 10^{-5}$		$\Gamma_{313}$	$K_0^*(1430)^0 f_0(980), f_0 \rightarrow \pi\pi$	$(2.7 \pm 0.9) \times 10^{-6}$	
$\Gamma_{253}$	$K^0\pi^0$	$(9.9 \pm 0.5) \times 10^{-6}$		$\Gamma_{314}$	$K_2^*(1430)^0 f_0(980), f_0 \rightarrow \pi\pi$	$(8.6 \pm 2.0) \times 10^{-6}$	
$\Gamma_{254}$	$\eta' K^0$	$(6.6 \pm 0.4) \times 10^{-5}$	S=1.4	$\Gamma_{315}$	$K^+\bar{K}^-$	$(1.3 \pm 0.5) \times 10^{-7}$	
$\Gamma_{255}$	$\eta' K^*(892)^0$	$(2.8 \pm 0.6) \times 10^{-6}$		$\Gamma_{316}$	$K^0\bar{K}^0$	$(1.21 \pm 0.16) \times 10^{-6}$	
$\Gamma_{256}$	$\eta' K_0^*(1430)^0$	$(6.3 \pm 1.6) \times 10^{-6}$		$\Gamma_{317}$	$K^0 K^-\pi^+$	$(6.5 \pm 0.8) \times 10^{-6}$	
	<b>K or K* modes</b>			$\Gamma_{318}$	$K^*(892)^\pm K^\mp$	$< 4 \times 10^{-7}$	CL=90%
				$\Gamma_{319}$	$\bar{K}^{*0} K^0 + K^{*0} \bar{K}^0$	$< 9.6 \times 10^{-7}$	CL=90%
				$\Gamma_{320}$	$K^+ K^-\pi^0$	$(2.2 \pm 0.6) \times 10^{-6}$	
				$\Gamma_{321}$	$K_S^0 K_S^0\pi^0$	$< 9 \times 10^{-7}$	CL=90%
				$\Gamma_{322}$	$K_S^0 K_S^0\eta$	$< 1.0 \times 10^{-6}$	CL=90%
				$\Gamma_{323}$	$K_S^0 K_S^0\eta'$	$< 2.0 \times 10^{-6}$	CL=90%

$\Gamma_{324}$	$K^0 K^+ K^-$	$(2.49 \pm 0.31) \times 10^{-5}$	S=3.0	$\Gamma_{393}$	$\omega \eta$	$(9.4 \pm 4.0 \pm 3.1) \times 10^{-7}$	
$\Gamma_{325}$	$K^0 \phi$	$(7.3 \pm 0.7) \times 10^{-6}$		$\Gamma_{394}$	$\omega \eta'$	$(1.0 \pm 0.5 \pm 0.4) \times 10^{-6}$	
$\Gamma_{326}$	$f_0(980) K^0, f_0 \rightarrow K^+ K^-$	$(7.0 \pm 3.5 \pm 3.0) \times 10^{-6}$		$\Gamma_{395}$	$\omega \rho^0$	$< 1.6$	CL=90%
$\Gamma_{327}$	$f_0(1500) K^0$	$(1.3 \pm 0.7 \pm 0.5) \times 10^{-5}$		$\Gamma_{396}$	$\omega f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 1.5$	CL=90%
$\Gamma_{328}$	$f'_2(1525)^0 K^0$	$(3 \pm 5 \pm 4) \times 10^{-7}$		$\Gamma_{397}$	$\omega \omega$	$(1.2 \pm 0.4) \times 10^{-6}$	
$\Gamma_{329}$	$f_0(1710) K^0, f_0 \rightarrow K^+ K^-$	$(4.4 \pm 0.9) \times 10^{-6}$		$\Gamma_{398}$	$\phi \pi^0$	$< 1.5$	CL=90%
$\Gamma_{330}$	$K^0 K^+ K^-$ nonresonant	$(3.3 \pm 1.0) \times 10^{-5}$		$\Gamma_{399}$	$\phi \eta$	$< 5$	CL=90%
$\Gamma_{331}$	$K_S^0 K_S^0 K_S^0$	$(6.0 \pm 0.5) \times 10^{-6}$	S=1.1	$\Gamma_{400}$	$\phi \eta'$	$< 5$	CL=90%
$\Gamma_{332}$	$f_0(980) K^0, f_0 \rightarrow K_S^0 K_S^0$	$(2.7 \pm 1.8) \times 10^{-6}$		$\Gamma_{401}$	$\phi \rho^0$	$< 3.3$	CL=90%
$\Gamma_{333}$	$f_0(1710) K^0, f_0 \rightarrow K_S^0 K_S^0$	$(5.0 \pm 5.0 \pm 2.6) \times 10^{-7}$		$\Gamma_{402}$	$\phi f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 3.8$	CL=90%
$\Gamma_{334}$	$f_0(2010) K^0, f_0 \rightarrow K_S^0 K_S^0$	$(5 \pm 6) \times 10^{-7}$		$\Gamma_{403}$	$\phi \omega$	$< 7$	CL=90%
$\Gamma_{335}$	$K_S^0 K_S^0 K_S^0$ nonresonant	$(1.33 \pm 0.31) \times 10^{-5}$		$\Gamma_{404}$	$\phi \phi$	$< 2.8$	CL=90%
$\Gamma_{336}$	$K_S^0 K_S^0 K_L^0$	$< 1.6$	CL=90%	$\Gamma_{405}$	$a_0(980) \pi^\pm \pi^\mp, a_0^\pm \rightarrow \eta \pi^\pm$	$< 3.1$	CL=90%
$\Gamma_{337}$	$K^*(892)^0 K^+ K^-$	$(2.75 \pm 0.26) \times 10^{-5}$		$\Gamma_{406}$	$a_0(1450) \pi^\pm \pi^\mp, a_0^\pm \rightarrow \eta \pi^\pm$	$< 2.3$	CL=90%
$\Gamma_{338}$	$K^*(892)^0 \phi$	$(1.00 \pm 0.05) \times 10^{-5}$		$\Gamma_{407}$	$\pi^+ \pi^- \pi^0$	$< 7.2$	CL=90%
$\Gamma_{339}$	$K^+ K^- \pi^+ \pi^-$ nonresonant	$< 7.17$	CL=90%	$\Gamma_{408}$	$\rho^0 \pi^0$	$(2.0 \pm 0.5) \times 10^{-6}$	
$\Gamma_{340}$	$K^*(892)^0 K^- \pi^+$	$(4.5 \pm 1.3) \times 10^{-6}$		$\Gamma_{409}$	$\rho^\mp \pi^\pm$	[h] $(2.30 \pm 0.23) \times 10^{-5}$	
$\Gamma_{341}$	$K^*(892)^0 \bar{K}^*(892)^0$	$(8 \pm 5) \times 10^{-7}$	S=2.2	$\Gamma_{410}$	$\pi^+ \pi^- \pi^+ \pi^-$	$< 1.12$	CL=90%
$\Gamma_{342}$	$K^+ K^- \pi^+ \pi^-$ nonresonant	$< 6.0$	CL=90%	$\Gamma_{411}$	$\rho^0 \pi^+ \pi^-$	$< 8.8$	CL=90%
$\Gamma_{343}$	$K^*(892)^0 K^+ \pi^-$	$< 2.2$	CL=90%	$\Gamma_{412}$	$\rho^0 \rho^0$	$(9.6 \pm 1.5) \times 10^{-7}$	
$\Gamma_{344}$	$K^*(892)^0 K^*(892)^0$	$< 2$	CL=90%	$\Gamma_{413}$	$f_0(980) \pi^+ \pi^-, f_0 \rightarrow \pi^+ \pi^-$	$< 3.0$	CL=90%
$\Gamma_{345}$	$K^*(892)^+ K^*(892)^-$	$< 2.0$	CL=90%	$\Gamma_{414}$	$\rho^0 f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$(7.8 \pm 2.5) \times 10^{-7}$	
$\Gamma_{346}$	$K_1(1400)^0 \phi$	$< 5.0$	CL=90%	$\Gamma_{415}$	$f_0(980) f_0(980), f_0 \rightarrow \pi^+ \pi^-, f_0 \rightarrow \pi^+ \pi^-$	$< 1.9$	CL=90%
$\Gamma_{347}$	$\phi(K\pi)_0^{*0} (1.60 < m_{K\pi} < 2.15)$	$(4.3 \pm 0.4) \times 10^{-6}$		$\Gamma_{416}$	$f_0(980) f_0(980), f_0 \rightarrow \pi^+ \pi^-, f_0 \rightarrow K^+ K^-$	$< 2.3$	CL=90%
$\Gamma_{348}$	$\phi(K\pi)_0^{*0} (1.60 < m_{K\pi} < 2.15)$	[f] $< 1.7$	CL=90%	$\Gamma_{417}$	$a_1(1260)^\mp \pi^\pm$	[h] $(2.6 \pm 0.5) \times 10^{-5}$	S=1.9
$\Gamma_{349}$	$K_0^*(1430)^0 K^- \pi^+$	$< 3.18$	CL=90%	$\Gamma_{418}$	$a_2(1320)^\mp \pi^\pm$	[h] $< 6.3$	CL=90%
$\Gamma_{350}$	$K_0^*(1430)^0 \bar{K}^*(892)^0$	$< 3.3$	CL=90%	$\Gamma_{419}$	$\pi^+ \pi^- \pi^0 \pi^0$	$< 3.1$	CL=90%
$\Gamma_{351}$	$K_0^*(1430)^0 \bar{K}_0^*(1430)^0$	$< 8.4$	CL=90%	$\Gamma_{420}$	$\rho^+ \rho^-$	$(2.77 \pm 0.19) \times 10^{-5}$	
$\Gamma_{352}$	$K_0^*(1430)^0 \phi$	$(3.9 \pm 0.8) \times 10^{-6}$		$\Gamma_{421}$	$a_1(1260)^0 \pi^0$	$< 1.1$	CL=90%
$\Gamma_{353}$	$K_0^*(1430)^0 K^*(892)^0$	$< 1.7$	CL=90%	$\Gamma_{422}$	$\omega \pi^0$	$< 5$	CL=90%
$\Gamma_{354}$	$K_0^*(1430)^0 K_0^*(1430)^0$	$< 4.7$	CL=90%	$\Gamma_{423}$	$\pi^+ \pi^+ \pi^- \pi^- \pi^0$	$< 9.0$	CL=90%
$\Gamma_{355}$	$K^*(1680)^0 \phi$	$< 3.5$	CL=90%	$\Gamma_{424}$	$a_1(1260)^\pm \rho^-$	$< 6.1$	CL=90%
$\Gamma_{356}$	$K^*(1780)^0 \phi$	$< 2.7$	CL=90%	$\Gamma_{425}$	$a_1(1260)^0 \rho^0$	$< 2.4$	CL=90%
$\Gamma_{357}$	$K^*(2045)^0 \phi$	$< 1.53$	CL=90%	$\Gamma_{426}$	$b_1^\mp \pi^\pm, b_1^\mp \rightarrow \omega \pi^\mp$	$(1.09 \pm 0.15) \times 10^{-5}$	
$\Gamma_{358}$	$K_2^*(1430)^0 \rho^0$	$< 1.1$	CL=90%	$\Gamma_{427}$	$b_1^0 \pi^0, b_1^0 \rightarrow \omega \pi^0$	$< 1.9$	CL=90%
$\Gamma_{359}$	$K_2^*(1430)^0 \phi$	$(6.8 \pm 0.9) \times 10^{-6}$	S=1.2	$\Gamma_{428}$	$b_1^\pm \rho^\pm, b_1^\pm \rightarrow \omega \pi^\pm$	$< 1.4$	CL=90%
$\Gamma_{360}$	$K^0 \phi \phi$	$(4.5 \pm 0.9) \times 10^{-6}$		$\Gamma_{429}$	$b_1^0 \rho^0, b_1^0 \rightarrow \omega \pi^0$	$< 3.4$	CL=90%
$\Gamma_{361}$	$\eta' \eta' K^0$	$< 3.1$	CL=90%	$\Gamma_{430}$	$\pi^+ \pi^+ \pi^+ \pi^- \pi^- \pi^-$	$< 3.0$	CL=90%
$\Gamma_{362}$	$\eta K^0 \gamma$	$(7.6 \pm 1.8) \times 10^{-6}$		$\Gamma_{431}$	$a_1(1260)^+ a_1(1260)^-, a_1^\pm \rightarrow 2\pi^+ \pi^-, a_1^\pm \rightarrow 2\pi^- \pi^+$	$(1.18 \pm 0.31) \times 10^{-5}$	
$\Gamma_{363}$	$\eta' K^0 \gamma$	$< 6.4$	CL=90%	$\Gamma_{432}$	$\pi^+ \pi^+ \pi^+ \pi^- \pi^- \pi^- \pi^0$	$< 1.1$	% CL=90%
$\Gamma_{364}$	$K^0 \phi \gamma$	$(2.7 \pm 0.7) \times 10^{-6}$		<b>Baryon modes</b>			
$\Gamma_{365}$	$K^+ \pi^- \gamma$	$(4.6 \pm 1.4) \times 10^{-6}$		$\Gamma_{433}$	$p \bar{p}$	$(1.5 \pm 0.7 \pm 0.5) \times 10^{-8}$	
$\Gamma_{366}$	$K^*(892)^0 \gamma$	$(4.33 \pm 0.15) \times 10^{-5}$		$\Gamma_{434}$	$p \bar{p} \pi^+ \pi^-$	$< 2.5$	CL=90%
$\Gamma_{367}$	$K^*(1410) \gamma$	$< 1.3$	CL=90%	$\Gamma_{435}$	$p \bar{p} K^0$	$(2.66 \pm 0.32) \times 10^{-6}$	
$\Gamma_{368}$	$K^+ \pi^- \gamma$ nonresonant	$< 2.6$	CL=90%	$\Gamma_{436}$	$\Theta(1540)^+ \bar{p}, \Theta^+ \rightarrow p K_S^0$	[i] $< 5$	CL=90%
$\Gamma_{369}$	$K^*(892)^0 X(214), X \rightarrow \mu^+ \mu^-$	[g] $< 2.26$	CL=90%	$\Gamma_{437}$	$f_J(2220) K^0, f_J \rightarrow p \bar{p}$	$< 4.5$	CL=90%
$\Gamma_{370}$	$K^0 \pi^+ \pi^- \gamma$	$(1.95 \pm 0.22) \times 10^{-5}$		$\Gamma_{438}$	$p \bar{p} K^*(892)^0$	$(1.24 \pm 0.28 \pm 0.25) \times 10^{-6}$	
$\Gamma_{371}$	$K^+ \pi^- \pi^0 \gamma$	$(4.1 \pm 0.4) \times 10^{-5}$		$\Gamma_{439}$	$f_J(2220) K_0^*, f_J \rightarrow p \bar{p}$	$< 1.5$	CL=90%
$\Gamma_{372}$	$K_1(1270)^0 \gamma$	$< 5.8$	CL=90%	$\Gamma_{440}$	$p \bar{\Lambda} \pi^-$	$(3.14 \pm 0.29) \times 10^{-6}$	
$\Gamma_{373}$	$K_1(1400)^0 \gamma$	$< 1.2$	CL=90%	$\Gamma_{441}$	$p \bar{\Lambda} \pi^- \gamma$	$< 6.5$	CL=90%
$\Gamma_{374}$	$K_2^*(1430)^0 \gamma$	$(1.24 \pm 0.24) \times 10^{-5}$		$\Gamma_{442}$	$p \bar{\Sigma}^-(1385)^-$	$< 2.6$	CL=90%
$\Gamma_{375}$	$K^*(1680)^0 \gamma$	$< 2.0$	CL=90%	$\Gamma_{443}$	$\Delta^0 \bar{\Lambda}$	$< 9.3$	CL=90%
$\Gamma_{376}$	$K_2^*(1780)^0 \gamma$	$< 8.3$	CL=90%	$\Gamma_{444}$	$p \bar{\Lambda} K^-$	$< 8.2$	CL=90%
$\Gamma_{377}$	$K_4^*(2045)^0 \gamma$	$< 4.3$	CL=90%	$\Gamma_{445}$	$p \bar{\Lambda} D^-$	$(2.5 \pm 0.4) \times 10^{-5}$	
<b>Light unflavored meson modes</b>				$\Gamma_{446}$	$p \bar{\Lambda} D^{*-}$	$(3.4 \pm 0.8) \times 10^{-5}$	
$\Gamma_{378}$	$\rho^0 \gamma$	$(8.6 \pm 1.5) \times 10^{-7}$		$\Gamma_{447}$	$p \bar{\Sigma}^0 \pi^-$	$< 3.8$	CL=90%
$\Gamma_{379}$	$\rho^0 X(214), X \rightarrow \mu^+ \mu^-$	[g] $< 1.73$	CL=90%	$\Gamma_{448}$	$\bar{\Lambda} \Lambda$	$< 3.2$	CL=90%
$\Gamma_{380}$	$\omega \gamma$	$(4.4 \pm 1.8 \pm 1.6) \times 10^{-7}$		$\Gamma_{449}$	$\bar{\Lambda} \Lambda K^0$	$(4.8 \pm 1.0 \pm 0.9) \times 10^{-6}$	
$\Gamma_{381}$	$\phi \gamma$	$< 8.5$	CL=90%	$\Gamma_{450}$	$\bar{\Lambda} \Lambda K^{*0}$	$(2.5 \pm 0.9 \pm 0.8) \times 10^{-6}$	
$\Gamma_{382}$	$\pi^+ \pi^-$	$(5.12 \pm 0.19) \times 10^{-6}$		$\Gamma_{451}$	$\bar{\Lambda} \Lambda D^0$	$(1.00 \pm 0.30 \pm 0.26) \times 10^{-5}$	
$\Gamma_{383}$	$\pi^0 \pi^0$	$(1.91 \pm 0.22) \times 10^{-6}$		$\Gamma_{452}$	$D^0 \Sigma^0 \bar{\Lambda} + c.c.$	$< 3.1$	CL=90%
$\Gamma_{384}$	$\eta \pi^0$	$(4.1 \pm 1.7) \times 10^{-7}$		$\Gamma_{453}$	$\Delta^0 \bar{\Sigma}^0$	$< 1.5$	CL=90%
$\Gamma_{385}$	$\eta \eta$	$< 1.0$	CL=90%	$\Gamma_{454}$	$\Delta^{++} \bar{\Delta}^{--}$	$< 1.1$	CL=90%
$\Gamma_{386}$	$\eta' \pi^0$	$(1.2 \pm 0.6) \times 10^{-6}$	S=1.7	$\Gamma_{455}$	$\bar{D}^0 p \bar{p}$	$(1.04 \pm 0.07) \times 10^{-4}$	
$\Gamma_{387}$	$\eta' \eta'$	$< 1.7$	CL=90%	$\Gamma_{456}$	$D_s^- \bar{\Lambda} p$	$(2.8 \pm 0.9) \times 10^{-5}$	
$\Gamma_{388}$	$\eta' \eta$	$< 1.2$	CL=90%	$\Gamma_{457}$	$\bar{D}^*(2007)^0 p \bar{p}$	$(9.9 \pm 1.1) \times 10^{-5}$	
$\Gamma_{389}$	$\eta' \rho^0$	$< 1.3$	CL=90%				
$\Gamma_{390}$	$\eta' f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 9$	CL=90%				
$\Gamma_{391}$	$\eta \rho^0$	$< 1.5$	CL=90%				
$\Gamma_{392}$	$\eta f_0(980), f_0 \rightarrow \pi^+ \pi^-$	$< 4$	CL=90%				

$\Gamma_{521}$	$K^*(892)^0 e^\pm \mu^\mp$	$LF$	$< 5.8$	$\times 10^{-7}$	CL=90%
$\Gamma_{522}$	$e^\pm \tau^\mp$	$LF$	$[h] < 2.8$	$\times 10^{-5}$	CL=90%
$\Gamma_{523}$	$\mu^\pm \tau^\mp$	$LF$	$[h] < 2.2$	$\times 10^{-5}$	CL=90%
$\Gamma_{524}$	invisible	$B1$	$< 2.4$	$\times 10^{-5}$	CL=90%
$\Gamma_{525}$	$\nu \bar{\nu} \gamma$	$B1$	$< 1.7$	$\times 10^{-5}$	CL=90%
$\Gamma_{526}$	$\Lambda_c^+ \mu^-$	$L,B$	$< 1.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{527}$	$\Lambda_c^+ e^-$	$L,B$	$< 4$	$\times 10^{-6}$	CL=90%

x <sub>7</sub>	47									
x <sub>34</sub>	0	0								
x <sub>46</sub>	0	0	43							
x <sub>72</sub>	0	0	6	13						
x <sub>123</sub>	0	0	10	4	1					
x <sub>183</sub>	0	0	0	0	0	0				
x <sub>185</sub>	0	0	0	0	0	0	0			
x <sub>236</sub>	0	0	0	0	0	0	0	0		
x <sub>241</sub>	0	0	0	0	0	0	0	0	19	
x <sub>245</sub>	0	0	0	0	0	0	0	6	0	0
x <sub>249</sub>	0	0	0	0	0	0	0	29	0	0
x <sub>252</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>286</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>317</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>324</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>338</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>382</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>412</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>506</sub>	0	0	0	0	0	0	4	0	0	0
x <sub>511</sub>	0	0	0	0	0	0	0	27	0	0

x <sub>6</sub>	x <sub>7</sub>	x <sub>34</sub>	x <sub>46</sub>	x <sub>72</sub>	x <sub>123</sub>	x <sub>183</sub>	x <sub>185</sub>	x <sub>236</sub>	x <sub>241</sub>	
x <sub>249</sub>	22									
x <sub>252</sub>	0	0								
x <sub>286</sub>	0	0	0							
x <sub>317</sub>	0	0	0	24						
x <sub>324</sub>	0	0	0	11	3					
x <sub>338</sub>	0	0	0	0	0	0				
x <sub>382</sub>	0	0	27	0	0	0	0			
x <sub>412</sub>	0	0	0	0	0	0	20	0		
x <sub>506</sub>	0	0	0	0	0	0	0	0	0	
x <sub>511</sub>	2	8	0	0	0	0	0	0	0	0

See key on page 601

## Meson Particle Listings

 $B^0$  $B^0$  BRANCHING RATIOS

For branching ratios in which the charge of the decaying  $B$  is not determined, see the  $B^\pm$  section.

 $\Gamma(e^+ \nu_e \text{ anything})/\Gamma_{\text{total}}$  $\Gamma_1/\Gamma$ 

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>10.33 ± 0.28 OUR EVALUATION</b>			
<b>10.14 ± 0.30 OUR AVERAGE</b>	Error includes scale factor of 1.1.		
10.46 ± 0.30 ± 0.23	<sup>1</sup> URQUIJO	07 BELL	$e^+ e^- \rightarrow \gamma(4S)$
9.64 ± 0.27 ± 0.33	<sup>2</sup> AUBERT, B	06Y BABR	$e^+ e^- \rightarrow \gamma(4S)$
10.78 ± 0.60 ± 0.69	<sup>3</sup> ARTUSO	97 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
9.3 ± 1.1 ± 1.5	ALBRECHT	94 ARG	$e^+ e^- \rightarrow \gamma(4S)$
9.9 ± 3.0 ± 0.9	HENDERSON	92 CLEO	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
10.32 ± 0.36 ± 0.35	<sup>4</sup> OKABE	05 BELL	Repl. by URQUIJO 07
10.9 ± 0.7 ± 1.1	ATHANAS	94 CLE2	Sup. by ARTUSO 97

<sup>1</sup> URQUIJO 07 report a measurement of  $(9.80 \pm 0.29 \pm 0.21)\%$  for the partial branching fraction of  $B \rightarrow e \nu_e X_c$  decay with electron energy above 0.6 GeV. We converted the result to  $B \rightarrow e \nu_e X$  branching fraction.

<sup>2</sup> The measurements are obtained for charged and neutral  $B$  mesons partial rates of semileptonic decay to electrons with momentum above 0.6 GeV/c in the  $B$  rest frame. The best precision on the ratio is achieved for a momentum threshold of 1.0 GeV:  $B(B^+ \rightarrow e^+ \nu_e X) / B(B^0 \rightarrow e^+ \nu_e X) = 1.074 \pm 0.041 \pm 0.026$ .

<sup>3</sup> ARTUSO 97 uses partial reconstruction of  $B \rightarrow D^* \ell \nu_\ell$  and inclusive semileptonic branching ratio from BARISH 96B ( $0.1049 \pm 0.0017 \pm 0.0043$ ).

<sup>4</sup> The measurements are obtained for charged and neutral  $B$  mesons partial rates of semileptonic decay to electrons with momentum above 0.6 GeV/c in the  $B$  rest frame, and their ratio of  $B(B^+ \rightarrow e^+ \nu_e X) / B(B^0 \rightarrow e^+ \nu_e X) = 1.08 \pm 0.05 \pm 0.02$ .

 $\Gamma(e^+ \nu_e X_c)/\Gamma_{\text{total}}$  $\Gamma_2/\Gamma$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>10.08 ± 0.30 ± 0.22</b>	<sup>1</sup> URQUIJO	07 BELL	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Measure the independent $B^+$ and $B^0$ partial branching fractions with electron threshold energies of 0.4 GeV.			

 $\Gamma(D^- \ell^+ \nu_\ell)/\Gamma_{\text{total}}$  $\Gamma_4/\Gamma$ 

$\ell$  denotes  $e$  or  $\mu$ , not the sum.

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0219 ± 0.0012 OUR EVALUATION</b>			
<b>0.0225 ± 0.0008 OUR AVERAGE</b>			
0.0231 ± 0.0003 ± 0.0011	<sup>1</sup> GLATTAUER	16 BELL	$e^+ e^- \rightarrow \gamma(4S)$
0.0221 ± 0.0011 ± 0.0011	<sup>2</sup> AUBERT	10 BABR	$e^+ e^- \rightarrow \gamma(4S)$
0.0209 ± 0.0013 ± 0.0018	<sup>3</sup> BARTELT	99 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
0.0235 ± 0.0020 ± 0.0044	<sup>4</sup> BUSKULIC	97 ALEP	$e^+ e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.0221 ± 0.0011 ± 0.0012	<sup>2</sup> AUBERT	08Q BABR	Repl. by AUBERT 10
0.0213 ± 0.0012 ± 0.0039	ABE	02E BELL	Repl. by GLATTAUER 16
0.0187 ± 0.0015 ± 0.0032	<sup>5</sup> ATHANAS	97 CLE2	Repl. by BARTELT 99
0.018 ± 0.006 ± 0.003	<sup>6</sup> FULTON	91 CLEO	$e^+ e^- \rightarrow \gamma(4S)$
0.020 ± 0.007 ± 0.006	<sup>7</sup> ALBRECHT	89J ARG	$e^+ e^- \rightarrow \gamma(4S)$

<sup>1</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side while the other, on the signal side, is partially reconstructed from  $B \rightarrow D \ell \nu$ .

<sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

<sup>4</sup> BUSKULIC 97 assumes fraction  $(B^+) = \text{fraction}(B^0) = (37.8 \pm 2.2)\%$  and PDG 96 values for  $B$  lifetime and branching ratio of  $D^*$  and  $D$  decays.

<sup>5</sup> ATHANAS 97 uses missing energy and missing momentum to reconstruct neutrino.

<sup>6</sup> FULTON 91 assumes assuming equal production of  $B^0$  and  $B^+$  at the  $\gamma(4S)$  and uses Mark III  $D$  and  $D^*$  branching ratios.

<sup>7</sup> ALBRECHT 89J reports  $0.018 \pm 0.006 \pm 0.005$ . We rescale using the method described in STONE 94 but with the updated PDG 94  $B(D^0 \rightarrow K^- \pi^+)$ .

 $\Gamma(D^- \ell^+ \nu_\ell)/\Gamma(e^+ \nu_e \text{ anything})$  $\Gamma_4/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.230 ± 0.011 ± 0.011</b>	<sup>1</sup> AUBERT	10 BABR	$e^+ e^- \rightarrow \gamma(4S)$

<sup>1</sup> Uses a fully reconstructed  $B$  meson on the recoil side.

 $\Gamma(D^- \ell^+ \nu_\ell)/\Gamma(D \ell^+ \nu_\ell \text{ anything})$  $\Gamma_4/\Gamma_3$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.215 ± 0.016 ± 0.013</b>	<sup>1</sup> AUBERT	07AN BABR	$e^+ e^- \rightarrow \gamma(4S)$

<sup>1</sup> Uses a fully reconstructed  $B$  meson on the recoil side.

 $\Gamma(D^- \tau^+ \nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_5/\Gamma$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.04 ± 0.35 ± 0.18	<sup>1</sup> AUBERT	08N BABR	Repl. by AUBERT 09s
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.			

 $\Gamma(D^- \tau^+ \nu_\tau)/\Gamma(D^- \ell^+ \nu_\ell)$  $\Gamma_5/\Gamma_4$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.469 ± 0.084 ± 0.053</b>	<sup>1,2</sup> LEES	12D BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.489 ± 0.165 ± 0.069	<sup>1</sup> AUBERT	09s BABR	Repl. by LEES 12D
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.			
<sup>2</sup> Uses $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$ and $\tau^+ \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$ and $e^+$ or $\mu^+$ as $\ell^+$ .			

 $\Gamma(D^*(2010)^- \ell^+ \nu_\ell)/\Gamma_{\text{total}}$  $\Gamma_6/\Gamma$ 

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0493 ± 0.0011 OUR EVALUATION</b>				
<b>0.0510 ± 0.0023 OUR FIT</b>				Error includes scale factor of 1.6.
<b>0.0509 ± 0.0022 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
0.0458 ± 0.0003 ± 0.0026		<sup>1</sup> DUNGEL	10 BELL	$e^+ e^- \rightarrow \gamma(4S)$
0.0549 ± 0.0016 ± 0.0025		<sup>2</sup> AUBERT	08Q BABR	$e^+ e^- \rightarrow \gamma(4S)$
0.0469 ± 0.0004 ± 0.0034		<sup>3</sup> AUBERT	08R BABR	$e^+ e^- \rightarrow \gamma(4S)$
0.0590 ± 0.0022 ± 0.0050		<sup>4</sup> ABDALLAH	04D DLPH	$e^+ e^- \rightarrow Z^0$
0.0609 ± 0.0019 ± 0.0040		<sup>5</sup> ADAM	03 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
0.0470 ± 0.0013 ± 0.0036		<sup>6</sup> ABREU	01H DLPH	$e^+ e^- \rightarrow Z$
0.0526 ± 0.0020 ± 0.0046		<sup>7</sup> ABBIENDI	00Q OPAL	$e^+ e^- \rightarrow Z$
0.0553 ± 0.0026 ± 0.0052		<sup>8</sup> BUSKULIC	97 ALEP	$e^+ e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0490 ± 0.0007 ± 0.0036		<sup>4</sup> AUBERT	05E BABR	Repl. by AUBERT 08R
0.0539 ± 0.0011 ± 0.0034		<sup>9</sup> ABDALLAH	04D DLPH	$e^+ e^- \rightarrow Z^0$
0.0459 ± 0.0023 ± 0.0040		<sup>10</sup> ABE	02F BELL	Repl. by DUNGEL 10
0.0609 ± 0.0019 ± 0.0040		<sup>11</sup> BRIERE	02 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
0.0508 ± 0.0021 ± 0.0066		<sup>12</sup> ACKERSTAFF	97G OPAL	Repl. by ABBIENDI 00q
0.0552 ± 0.0017 ± 0.0068		<sup>13</sup> ABREU	96P DLPH	Repl. by ABREU 01H
0.0449 ± 0.0032 ± 0.0039	376	<sup>14</sup> BARISH	95 CLE2	Repl. by ADAM 03
0.0518 ± 0.0030 ± 0.0062	410	<sup>15</sup> BUSKULIC	95N ALEP	Sup. by BUSKULIC 97
0.045 ± 0.003 ± 0.004		<sup>16</sup> ALBRECHT	94 ARG	$e^+ e^- \rightarrow \gamma(4S)$
0.047 ± 0.005 ± 0.005	235	<sup>17</sup> ALBRECHT	93 ARG	$e^+ e^- \rightarrow \gamma(4S)$
seen 398		<sup>18</sup> SANGHERA	93 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
0.070 ± 0.018 ± 0.014		<sup>19</sup> ANTREASANYAN	90B CBAL	$e^+ e^- \rightarrow \gamma(4S)$
		<sup>20</sup> ALBRECHT	89C ARG	$e^+ e^- \rightarrow \gamma(4S)$
0.060 ± 0.010 ± 0.014		<sup>21</sup> ALBRECHT	89J ARG	$e^+ e^- \rightarrow \gamma(4S)$
0.040 ± 0.004 ± 0.006		<sup>22</sup> BORTOLETTO	89B CLEO	$e^+ e^- \rightarrow \gamma(4S)$
0.070 ± 0.012 ± 0.019	47	<sup>23</sup> ALBRECHT	87J ARG	$e^+ e^- \rightarrow \gamma(4S)$

<sup>1</sup> Uses fully reconstructed  $D^{*-} \ell^+ \nu$  events ( $\ell = e$  or  $\mu$ ).

<sup>2</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side.

<sup>3</sup> Measured using fully reconstructed  $D^*$  sample and a simultaneous fit to the Caprini-Lellouch-Neubert form factor parameters:  $\rho^2 = 1.191 \pm 0.048 \pm 0.028$ ,  $R_1(1) = 1.429 \pm 0.061 \pm 0.044$ , and  $R_2(1) = 0.827 \pm 0.038 \pm 0.022$ .

<sup>4</sup> Measured using fully reconstructed  $D^*$  sample.

<sup>5</sup> Uses the combined fit of both  $B^0 \rightarrow D^*(2010)^- \ell \nu$  and  $B^+ \rightarrow \bar{D}^*(2010)^0 \ell \nu$  samples.

<sup>6</sup> ABREU 01H measured using about 5000 partial reconstructed  $D^*$  sample.

<sup>7</sup> ABBIENDI 00Q assumes the fraction  $B(b \rightarrow B^0) = (39.7^{+1.8}_{-2.2})\%$ . This result is an average of two methods using exclusive and partial  $D^*$  reconstruction.

<sup>8</sup> BUSKULIC 97 assumes fraction  $(B^+) = \text{fraction}(B^0) = (37.8 \pm 2.2)\%$  and PDG 96 values for  $B$  lifetime and  $D^*$  and  $D$  branching fractions.

<sup>9</sup> Combines with previous partial reconstructed  $D^*$  measurement.

<sup>10</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

<sup>11</sup> The results are based on the same analysis and data sample reported in ADAM 03.

<sup>12</sup> ACKERSTAFF 97G assumes fraction  $(B^+) = \text{fraction}(B^0) = (37.8 \pm 2.2)\%$  and PDG 96 values for  $B$  lifetime and branching ratio of  $D^*$  and  $D$  decays.

<sup>13</sup> ABREU 96P result is the average of two methods using exclusive and partial  $D^*$  reconstruction.

<sup>14</sup> BARISH 95 use  $B(D^0 \rightarrow K^- \pi^+) = (3.91 \pm 0.08 \pm 0.17)\%$  and  $B(D^{*+} \rightarrow D^0 \pi^+) = (68.1 \pm 1.0 \pm 1.3)\%$ .

<sup>15</sup> BUSKULIC 95N assumes fraction  $(B^+) = \text{fraction}(B^0) = 38.2 \pm 1.3 \pm 2.2\%$  and  $\tau_{B^0} = 1.58 \pm 0.06$  ps.  $\Gamma(D^{*-} \ell^+ \nu_\ell)/\text{total} = [5.18 - 0.13(\text{fraction}(B^0) - 38.2) - 1.5(\tau_{B^0} - 1.58)]\%$ .

<sup>16</sup> ALBRECHT 94 assumes  $B(D^{*+} \rightarrow D^0 \pi^+) = 68.1 \pm 1.0 \pm 1.3\%$ . Uses partial reconstruction of  $D^{*+}$  and is independent of  $D^0$  branching ratios.

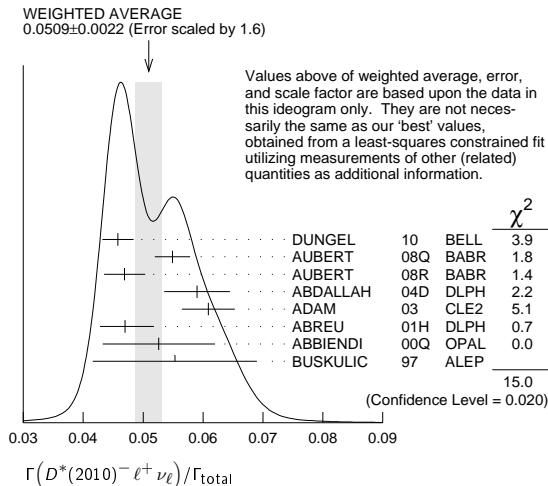
<sup>17</sup> ALBRECHT 93 reports  $0.052 \pm 0.005 \pm 0.006$ . We rescale using the method described in STONE 94 but with the updated PDG 94  $B(D^0 \rightarrow K^- \pi^+)$ . We have taken their average  $e$  and  $\mu$  value. They also obtain  $\alpha = 2s\Gamma^0/(\Gamma^- + \Gamma^+) - 1 = 1.1 \pm 0.4 \pm 0.2$ ,  $A_{AF} = 3/4 * (\Gamma^- - \Gamma^+)/\Gamma = 0.2 \pm 0.08 \pm 0.06$  and a value of  $|V_{cb}| = 0.036 - 0.045$  depending on model assumptions.

<sup>18</sup> Combining  $\bar{D}^{*0} \ell^+ \nu_\ell$  and  $\bar{D}^{*-} \ell^+ \nu_\ell$  SANGHERA 93 test  $V-A$  structure and fit the decay angular distributions to obtain  $A_{FB} = 3/4 * (\Gamma^- - \Gamma^+)/\Gamma = 0.14 \pm 0.06 \pm 0.03$ .

# Meson Particle Listings

$B^0$

- Assuming a value of  $V_{cb}$ , they measure  $V$ ,  $A_1$ , and  $A_2$ , the three form factors for the  $D^* \ell \nu_\ell$  decay, where results are slightly dependent on model assumptions.
- <sup>19</sup> ANTREASYAN 90b is average over  $B$  and  $\bar{D}^*$ (2010) charge states.
- <sup>20</sup> The measurement of ALBRECHT 89c suggests a  $D^*$  polarization  $\gamma_L/\gamma_T$  of  $0.85 \pm 0.45$ , or  $\alpha = 0.7 \pm 0.9$ .
- <sup>21</sup> ALBRECHT 89j is ALBRECHT 87j value rescaled using  $B(D^*(2010)^- \rightarrow D^0 \pi^-) = 0.57 \pm 0.04 \pm 0.04$ . Superseded by ALBRECHT 93.
- <sup>22</sup> We have taken average of the the BORTOLETTO 89b values for electrons and muons,  $0.046 \pm 0.005 \pm 0.007$ . We rescale using the method described in STONE 94 but with the updated PDG 94  $B(D^0 \rightarrow K^- \pi^+) = 0.65 \pm 0.66 \pm 0.25$ .
- <sup>23</sup> ALBRECHT 87j assume  $\mu$ -e universality, the  $B(\tau(4S) \rightarrow B^0 \bar{B}^0) = 0.45$ , the  $B(D^0 \rightarrow K^- \pi^+) = (0.042 \pm 0.004 \pm 0.004)$ , and the  $B(D^*(2010)^- \rightarrow D^0 \pi^-) = 0.49 \pm 0.08$ . Superseded by ALBRECHT 89j.



$\Gamma(D^*(2010)^- \ell^+ \nu_\ell) / \Gamma(D \ell^+ \nu_\ell \text{ anything})$				$\Gamma_6/\Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.537±0.031±0.036</b>	<sup>1</sup> AUBERT	07AN	BABR $e^+ e^- \rightarrow \tau(4S)$	
<sup>1</sup> Uses a fully reconstructed $B$ meson on the recoil side.				

$\Gamma(D^*(2010)^- \tau^+ \nu_\tau) / \Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.78±0.17 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>2.02<sup>+0.40</sup><sub>-0.37</sub>±0.37</b>	<sup>1</sup> MATYJA	07	BELL $e^+ e^- \rightarrow \tau(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.11±0.51±0.06	<sup>2</sup> AUBERT	08N	BABR Repl. by AUBERT 09s	
<sup>1</sup> Observed in the recoil of the accompanying $B$ meson.				
<sup>2</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				

$\Gamma(D^*(2010)^- \tau^+ \nu_\tau) / \Gamma(D^*(2010)^- \ell^+ \nu_\ell)$				$\Gamma_7/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.349±0.030 OUR FIT</b>				
<b>0.345±0.030 OUR AVERAGE</b>				
0.336±0.027±0.030	<sup>1</sup> AAIJ	15Q	LHCB $pp$ at 7, 8 TeV	
0.355±0.039±0.021	<sup>2,3</sup> LEES	12b	BABR $e^+ e^- \rightarrow \tau(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.207±0.095±0.008	<sup>2</sup> AUBERT	09s	BABR Repl. by LEES 12b	
<sup>1</sup> Uses $\tau^+ \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$ and $\mu^+$ as $\ell^+$ .				
<sup>2</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				
<sup>3</sup> Uses $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$ and $\tau^+ \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$ and $e^+$ or $\mu^+$ as $\ell^+$ .				

$\Gamma(\bar{D}^0 \pi^- \ell^+ \nu_\ell) / \Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>4.3±0.6 OUR AVERAGE</b>				
4.3±0.8±0.3	<sup>1</sup> AUBERT	08Q	BABR $e^+ e^- \rightarrow \tau(4S)$	
4.3±0.9±0.2	<sup>1,2</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.4±1.0±0.2	<sup>3</sup> LIVENTSEV	05	BELL Repl. by LIVENTSEV 08	
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				
<sup>2</sup> LIVENTSEV 08 reports $(4.2 \pm 0.7 \pm 0.6) \times 10^{-3}$ from a measurement of $[\Gamma(B^0 \rightarrow \bar{D}^0 \pi^- \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \ell^+ \nu_\ell)]$ assuming $B(B^0 \rightarrow D^- \ell^+ \nu_\ell) = (2.12 \pm 0.20) \times 10^{-2}$ , which we rescale to our best value $B(B^0 \rightarrow D^- \ell^+ \nu_\ell) = (2.19 \pm 0.12) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>3</sup> LIVENTSEV 05 reports $[\Gamma(B^0 \rightarrow \bar{D}^0 \pi^- \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell)] = 0.15 \pm 0.03 \pm 0.03$ which we multiply by our best value $B(B^+ \rightarrow \bar{D}^0 \ell^+ \nu_\ell) = (2.27 \pm 0.11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(D_0^{*0}(2400)^- \ell^+ \nu_\ell, D_0^{*-} \rightarrow \bar{D}^0 \pi^-) / \Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>3.0±1.2 OUR AVERAGE</b>	Error includes scale factor of 1.8.			
4.4±0.8±0.6	<sup>1</sup> AUBERT	08BL	BABR $e^+ e^- \rightarrow \tau(4S)$	
2.0±0.7±0.5	<sup>1</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$	
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				

$\Gamma(D_2^{*0}(2460)^- \ell^+ \nu_\ell, D_2^{*-} \rightarrow \bar{D}^0 \pi^-) / \Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.21±0.33 OUR AVERAGE</b>	Error includes scale factor of 1.8.			
1.10±0.17±0.08	<sup>1</sup> AUBERT	09Y	BABR $e^+ e^- \rightarrow \tau(4S)$	
2.2 ± 0.4 ± 0.4	<sup>2</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$	
<sup>1</sup> Uses a simultaneous fit of all $B$ semileptonic decays without full reconstruction of events. AUBERT 09y reports $B(B^0 \rightarrow \bar{D}_2^{*0}(2460)^- \ell^+ \nu_\ell) \cdot B(\bar{D}_2^{*0}(2460)^- \rightarrow \bar{D}^{*0} \pi^-) = (1.77 \pm 0.26 \pm 0.11) \times 10^{-3}$ and the authors have provided us the individual measurement.				
<sup>2</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				

$\Gamma(\bar{D}^{*0} n \pi \ell^+ \nu_\ell (n \geq 1)) / \Gamma(D \ell^+ \nu_\ell \text{ anything})$				$\Gamma_{11}/\Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.248±0.032±0.030</b>	<sup>1</sup> AUBERT	07AN	BABR $e^+ e^- \rightarrow \tau(4S)$	
<sup>1</sup> Uses a fully reconstructed $B$ meson on the recoil side.				

$\Gamma(\bar{D}^{*0} \pi^- \ell^+ \nu_\ell) / \Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>4.9±0.8 OUR AVERAGE</b>				
4.8±0.8±0.4	<sup>1</sup> AUBERT	08Q	BABR $e^+ e^- \rightarrow \tau(4S)$	
5.8±2.3±0.3	<sup>1,2</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.7±1.3±0.2	<sup>3,4</sup> LIVENTSEV	05	BELL Repl. by LIVENTSEV 08	
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				
<sup>2</sup> LIVENTSEV 08 reports $(5.6 \pm 2.1 \pm 0.8) \times 10^{-3}$ from a measurement of $[\Gamma(B^0 \rightarrow \bar{D}^{*0} \pi^- \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \ell^+ \nu_\ell)]$ assuming $B(B^0 \rightarrow D^- \ell^+ \nu_\ell) = (2.12 \pm 0.20) \times 10^{-2}$ , which we rescale to our best value $B(B^0 \rightarrow D^- \ell^+ \nu_\ell) = (2.19 \pm 0.12) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>3</sup> Excludes $D^*$ contribution to $D\pi$ modes.				
<sup>4</sup> LIVENTSEV 05 reports $[\Gamma(B^0 \rightarrow \bar{D}^{*0} \pi^- \ell^+ \nu_\ell) / \Gamma_{\text{total}}] / [B(B^+ \rightarrow \bar{D}^{*0}(2007)^0 \ell^+ \nu_\ell)] = 0.10 \pm 0.02 \pm 0.01$ which we multiply by our best value $B(B^+ \rightarrow \bar{D}^{*0}(2007)^0 \ell^+ \nu_\ell) = (5.69 \pm 0.19) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(D_1(2420)^- \ell^+ \nu_\ell, D_1^- \rightarrow \bar{D}^{*0} \pi^-) / \Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>2.80±0.28 OUR AVERAGE</b>				
2.78±0.24±0.25	<sup>1</sup> AUBERT	09Y	BABR $e^+ e^- \rightarrow \tau(4S)$	
2.7 ± 0.4 ± 0.3	<sup>2</sup> AUBERT	08BL	BABR $e^+ e^- \rightarrow \tau(4S)$	
5.4 ± 1.9 ± 0.9	<sup>2</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$	
<sup>1</sup> Uses a simultaneous measurement of all $B$ semileptonic decays without full reconstruction of events.				
<sup>2</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				

$\Gamma(D_1'(2430)^- \ell^+ \nu_\ell, D_1'^- \rightarrow \bar{D}^{*0} \pi^-) / \Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>3.1±0.7±0.5</b>		<sup>1</sup> AUBERT	08BL	BABR $e^+ e^- \rightarrow \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.0	90	<sup>1</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				

$\Gamma(D_2^{*0}(2460)^- \ell^+ \nu_\ell, D_2^{*-} \rightarrow \bar{D}^{*0} \pi^-) / \Gamma_{\text{total}}$				$\Gamma_{15}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.68±0.12 OUR AVERAGE</b>				
0.67±0.12±0.05		<sup>1</sup> AUBERT	09Y	BABR $e^+ e^- \rightarrow \tau(4S)$
0.7 ± 0.2 ± 0.2		<sup>2</sup> AUBERT	08BL	BABR $e^+ e^- \rightarrow \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0	90	<sup>2</sup> LIVENTSEV	08	BELL $e^+ e^- \rightarrow \tau(4S)$
<sup>1</sup> Uses a simultaneous fit of all $B$ semileptonic decays without full reconstruction of events. AUBERT 09y reports $B(B^0 \rightarrow \bar{D}_2^{*0}(2460)^- \ell^+ \nu_\ell) \cdot B(\bar{D}_2^{*0}(2460)^- \rightarrow \bar{D}^{*0} \pi^-) = (1.77 \pm 0.26 \pm 0.11) \times 10^{-3}$ and the authors have provided us the individual measurement.				
<sup>2</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.				

$\Gamma(D^- \pi^+ \pi^- \ell^+ \nu_\ell) / \Gamma(D^- \ell^+ \nu_\ell)$				$\Gamma_{16}/\Gamma_4$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>5.8±1.8±1.2</b>	<sup>1</sup> LEES	16	BABR $e^+ e^- \rightarrow \tau(4S)$	
<sup>1</sup> Measurement used electrons and muons as leptons.				

$\Gamma(D^{*-}\pi^+\pi^-\ell^+\nu_\ell)/\Gamma(D^{*}(2010)^-\ell^+\nu_\ell)$  $\Gamma_{17}/\Gamma_6$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.8 \pm 0.8 \pm 0.6</math></b>	<sup>1</sup> LEES	16	BABR $e^+e^- \rightarrow \Upsilon(4S)$

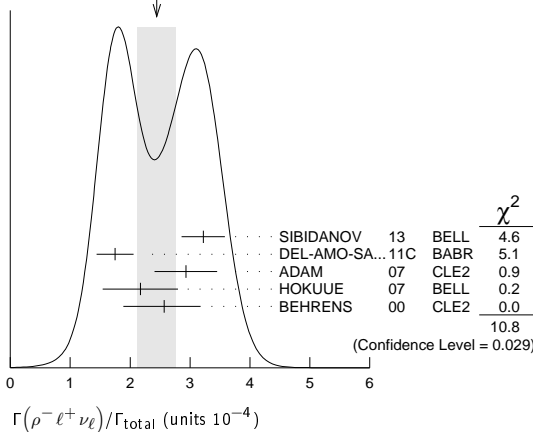
<sup>1</sup> Measurement used electrons and muons as leptons. $\Gamma(\rho^-\ell^+\nu_\ell)/\Gamma_{\text{total}}$  $\Gamma_{18}/\Gamma$  $\ell = e \text{ or } \mu$ , not sum over  $e$  and  $\mu$  modes.

“OUR EVALUATION” has been obtained by the Heavy Flavor Averaging Group (HFAG) by including both  $B^0$  and  $B^+$  decays. The average assumes equality of the semileptonic decay width for these isospin conjugate states.

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.94 \pm 0.11 \pm 0.18</math> OUR EVALUATION</b>				

 **$2.45 \pm 0.32$  OUR AVERAGE** Error includes scale factor of 1.6. See the ideogram below.

$3.22 \pm 0.27 \pm 0.24$	<sup>1</sup> SIBIDANOV	13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.75 \pm 0.15 \pm 0.27$	<sup>2</sup> DEL-AMO-SA...	11c	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$2.93 \pm 0.37 \pm 0.37$	<sup>3</sup> ADAM	07	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$2.17 \pm 0.54 \pm 0.32$	<sup>4</sup> HOKUUE	07	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$2.57 \pm 0.29 \pm 0.53$	<sup>5</sup> BEHRENS	00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$2.14 \pm 0.21 \pm 0.56$	<sup>2</sup> AUBERT,B	05o	BABR	Repl. by DEL-AMO-SANCHEZ 11c
$2.17 \pm 0.34 \pm 0.62$	<sup>6</sup> ATHAR	03	CLE2	Repl. by ADAM 07
$3.29 \pm 0.42 \pm 0.72$	<sup>7</sup> AUBERT	03E	BABR	Repl. by AUBERT,B 05o
$2.69 \pm 0.41 \pm 0.61$	<sup>8</sup> BEHRENS	00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$2.5 \pm 0.4 \pm 0.7$	<sup>9</sup> ALEXANDER	96T	CLE2	Repl. by BEHRENS 00
<4.1	<sup>10</sup> BEAN	93B	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> The signal events are tagged by a second  $B$  meson reconstructed in the fully hadronic decays.<sup>2</sup>  $B^+$  and  $B^0$  decays combined assuming isospin symmetry. Systematic errors include both experimental and form-factor uncertainties.<sup>3</sup> The  $B^0$  and  $B^+$  results are combined assuming the isospin,  $B$  lifetimes, and relative charged/neutral  $B$  production at the  $\Upsilon(4S)$ .<sup>4</sup> The signal events are tagged by a second  $B$  meson reconstructed in the semileptonic mode  $B \rightarrow D^{(*)}\ell\nu_\ell$ .<sup>5</sup> Averaging with ALEXANDER 96T results including experimental and theoretical correlations considered, BEHRENS 00 reports systematic errors  $+0.33 \pm 0.41$ , where the second error is theoretical model dependence. We combine these in quadrature.<sup>6</sup> ATHAR 03 reports systematic errors  $+0.47 \pm 0.41 \pm 0.01$ , which are experimental systematic, systematic due to residual form-factor uncertainties in the signal, and systematic due to residual form-factor uncertainties in the cross-feed modes, respectively. We combine these in quadrature.<sup>7</sup> Uses isospin constraints and extrapolation to all electron energies according to five different form-factor calculations. The second error combines the systematic and theoretical uncertainties in quadrature.<sup>8</sup> BEHRENS 00 reports  $+0.35 \pm 0.40 \pm 0.50$ , where the second error is the theoretical model dependence. We combine these in quadrature.  $B^+$  and  $B^0$  decays combined using isospin symmetry:  $\Gamma(B^0 \rightarrow \rho^-\ell^+\nu) = 2\Gamma(B^+ \rightarrow \rho^0\ell^+\nu) \approx 2\Gamma(B^+ \rightarrow \omega\ell^+\nu)$ . No evidence for  $\omega\ell\nu$  is reported.<sup>9</sup> ALEXANDER 96T reports  $+0.5 \pm 0.5$  where the second error is the theoretical model dependence. We combine these in quadrature.  $B^+$  and  $B^0$  decays combined using isospin symmetry:  $\Gamma(B^0 \rightarrow \rho^-\ell^+\nu) = 2\Gamma(B^+ \rightarrow \rho^0\ell^+\nu) \approx 2\Gamma(B^+ \rightarrow \omega\ell^+\nu)$ . No evidence for  $\omega\ell\nu$  is reported.<sup>10</sup> BEAN 93B limit set using ISGW Model. Using isospin and the quark model to combine  $\Gamma(\rho^0\ell^+\nu_\ell)$  and  $\Gamma(\omega\ell^+\nu_\ell)$  with this result, they obtain a limit  $<(1.6-2.7) \times 10^{-4}$  at 90% CL for  $B^+ \rightarrow (\omega \text{ or } \rho^0)\ell^+\nu_\ell$ . The range corresponds to the ISGW, WSB, and KS models. An upper limit on  $|V_{ub}/V_{cb}| < 0.08-0.13$  at 90% CL is derived as well.WEIGHTED AVERAGE  
 $2.45 \pm 0.32$  (Error scaled by 1.6) $\Gamma(\pi^-\ell^+\nu_\ell)/\Gamma_{\text{total}}$  $\Gamma_{19}/\Gamma$ 

“OUR EVALUATION” is provided by the Heavy Flavor Averaging Group (HFAG) and the procedure is described at <http://www.slac.stanford.edu/xorg/hfag/>.

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.45 \pm 0.05</math> OUR EVALUATION</b>			

 **$1.46 \pm 0.04$  OUR AVERAGE**

$1.49 \pm 0.09 \pm 0.07$	<sup>1</sup> SIBIDANOV	13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.47 \pm 0.05 \pm 0.06$	<sup>2,3</sup> LEES	12AA	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.41 \pm 0.05 \pm 0.07$	<sup>4</sup> DEL-AMO-SA...	11c	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.49 \pm 0.04 \pm 0.07$	<sup>2</sup> HA	11	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.54 \pm 0.17 \pm 0.09$	<sup>4</sup> AUBERT	08AV	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.37 \pm 0.15 \pm 0.11$	<sup>5,6</sup> ADAM	07	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$1.38 \pm 0.19 \pm 0.14$	<sup>7</sup> HOKUUE	07	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.42 \pm 0.05 \pm 0.08$	<sup>2</sup> DEL-AMO-SA...	11F	BABR	Repl. by LEES 12AA
$1.46 \pm 0.07 \pm 0.08$	<sup>8</sup> AUBERT	07J	BABR	Repl. by DEL-AMO-SANCHEZ 11F
$1.33 \pm 0.17 \pm 0.11$	<sup>9</sup> AUBERT,B	06K	BABR	Repl. by AUBERT 08AV
$1.38 \pm 0.10 \pm 0.18$	<sup>10</sup> AUBERT,B	05o	BABR	Repl. by DEL-AMO-SANCHEZ 11c
$1.33 \pm 0.18 \pm 0.13$	<sup>11</sup> ATHAR	03	CLE2	Repl. by ADAM 07
$1.8 \pm 0.4 \pm 0.4$	<sup>12</sup> ALEXANDER	96T	CLE2	Repl. by ATHAR 03

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>2</sup> Uses loose neutrino reconstruction technique. Assumes  $B(\Upsilon(4S) \rightarrow B^+\pi^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .<sup>3</sup> Reports also a branching fraction value  $B(B^0 \rightarrow \pi^-\ell^+\nu) = (1.45 \pm 0.04 \pm 0.06) \times 10^{-4}$  from the decays of  $B^+$  and  $B^0$  that are combined using the isospin symmetry relation.<sup>4</sup> Using the isospin symmetry relation,  $B^+$  and  $B^0$  branching fractions are combined.<sup>5</sup> The  $B^0$  and  $B^+$  results are combined assuming the isospin,  $B$  lifetimes, and relative charged/neutral  $B$  production at the  $\Upsilon(4S)$ .<sup>6</sup> Also report the rate for  $q^2 > 16 \text{ GeV}^2$  of  $(0.41 \pm 0.08 \pm 0.04) \times 10^{-4}$  from which they obtain  $|V_{ub}| = 3.6 \pm 0.4 \pm 0.2 \pm 0.6$  (last error is from theory).<sup>7</sup> The signal events are tagged by a second  $B$  meson reconstructed in the semileptonic mode  $B \rightarrow D^{(*)}\ell\nu_\ell$ .<sup>8</sup> The analysis uses events in which the signal  $B$  decays are reconstructed with an innovative loose neutrino reconstruction technique.<sup>9</sup> The signals are tagged by a second  $B$  meson reconstructed in a semileptonic or hadronic decay. The  $B^0$  and  $B^+$  results are combined assuming the isospin symmetry.<sup>10</sup>  $B^+$  and  $B^0$  decays combined assuming isospin symmetry. Systematic errors include both experimental and form-factor uncertainties.<sup>11</sup> ATHAR 03 reports systematic errors  $0.11 \pm 0.01 \pm 0.07$ , which are experimental systematic, systematic due to residual form-factor uncertainties in the signal, and systematic due to residual form-factor uncertainties in the cross-feed modes, respectively. We combine these in quadrature.<sup>12</sup> ALEXANDER 96T gives systematic errors  $\pm 0.3 \pm 0.2$  where the second error reflects the estimated model dependence. We combine these in quadrature. Assumes isospin symmetry:  $\Gamma(B^0 \rightarrow \pi^-\ell^+\nu) = 2 \times \Gamma(B^+ \rightarrow \pi^0\ell^+\nu)$ . $\Gamma(\pi^-\mu^+\nu_\mu)/\Gamma_{\text{total}}$  $\Gamma_{20}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
seen	<sup>1</sup> ALBRECHT	91c	ARG

<sup>1</sup> In ALBRECHT 91c, one event is fully reconstructed providing evidence for the  $b \rightarrow u$  transition. $\Gamma(\pi^-\tau^+\nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{21}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.5 \times 10^{-4}</math></b>	90	<sup>1</sup> HAMER	16	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(K^+\text{anything})/\Gamma_{\text{total}}$  $\Gamma_{22}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.78 \pm 0.08</math></b>	<sup>1</sup> ALBRECHT	96D	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Average multiplicity. $\Gamma(D^0X)/\Gamma_{\text{total}}$  $\Gamma_{23}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.081 \pm 0.014 \pm 0.005</math></b>	<sup>1</sup> AUBERT	07N	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.063 \pm 0.019 \pm 0.005$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
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<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties. $\Gamma(\bar{D}^0X)/\Gamma_{\text{total}}$  $\Gamma_{24}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.474 \pm 0.020 \pm 0.020</math></b>	<sup>1</sup> AUBERT	07N	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.511 \pm 0.031 \pm 0.028$	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
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<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.

# Meson Particle Listings

$B^0$

$\Gamma(D^0 X)/[\Gamma(D^0 X) + \Gamma(\overline{D}^0 X)]$	$\Gamma_{23}/(\Gamma_{23} + \Gamma_{24})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.146 ± 0.022 ± 0.006</b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.110 ± 0.031 ± 0.008	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(D^+ X)/\Gamma_{\text{total}}$	$\Gamma_{25}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.039</b>	90	<sup>1</sup> AUBERT	07N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.051	90	<sup>1</sup> AUBERT,BE	04B	BABR Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D^- X)/\Gamma_{\text{total}}$	$\Gamma_{26}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.369 ± 0.016 ± 0.030 ± 0.027</b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.397 ± 0.030 ± 0.040 ± 0.038	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D^+ X)/[\Gamma(D^+ X) + \Gamma(D^- X)]$	$\Gamma_{25}/(\Gamma_{25} + \Gamma_{26})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.058 ± 0.028 ± 0.006</b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.055 ± 0.040 ± 0.006	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(D_s^+ X)/\Gamma_{\text{total}}$	$\Gamma_{27}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.103 ± 0.012 ± 0.017 ± 0.014</b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.109 ± 0.021 ± 0.039 ± 0.024	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D_s^- X)/\Gamma_{\text{total}}$	$\Gamma_{28}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.026</b>	90	<sup>1</sup> AUBERT	07N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.087	90	<sup>1</sup> AUBERT,BE	04B	BABR Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(D_s^+ X)/[\Gamma(D_s^+ X) + \Gamma(D_s^- X)]$	$\Gamma_{27}/(\Gamma_{27} + \Gamma_{28})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.879 ± 0.066 ± 0.005</b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.733 ± 0.092 ± 0.010	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(A_c^+ X)/\Gamma_{\text{total}}$	$\Gamma_{29}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.031</b>	90	<sup>1</sup> AUBERT	07N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.038	90	<sup>1</sup> AUBERT,BE	04B	BABR Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(\overline{A}_c^- X)/\Gamma_{\text{total}}$	$\Gamma_{30}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.05 ± 0.010 ± 0.019 ± 0.011</b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.049 ± 0.017 ± 0.018 ± 0.011	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N
<sup>1</sup> Events are selected by completely reconstructing one $B$ and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.				

$\Gamma(A_c^+ X)/[\Gamma(A_c^+ X) + \Gamma(\overline{A}_c^- X)]$	$\Gamma_{29}/(\Gamma_{29} + \Gamma_{30})$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.243 ± 0.119 ± 0.121 ± 0.003</b>	AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.286 ± 0.142 ± 0.007	AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

$\Gamma(\Upsilon X)/\Gamma_{\text{total}}$	$\Gamma_{31}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.947 ± 0.030 ± 0.045 ± 0.040</b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.039 ± 0.051 ± 0.063 ± 0.058	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.

$\Gamma(c X)/\Gamma_{\text{total}}$	$\Gamma_{32}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.246 ± 0.024 ± 0.021 ± 0.017</b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.237 ± 0.036 ± 0.041 ± 0.027	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.

$\Gamma(\overline{c} c X)/\Gamma_{\text{total}}$	$\Gamma_{33}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>1.193 ± 0.030 ± 0.053 ± 0.049</b>	<sup>1</sup> AUBERT	07N	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.276 ± 0.062 ± 0.088 ± 0.074	<sup>1</sup> AUBERT,BE	04B	BABR	Repl. by AUBERT 07N

<sup>1</sup> Events are selected by completely reconstructing one  $B$  and searching for a reconstructed charmed particle in the rest of the event. The last error includes systematic and charm branching ratio uncertainties.

$\Gamma(D^- \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{34}/\Gamma$			
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.52 ± 0.13 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>2.68 ± 0.13 OUR AVERAGE</b>				
2.55 ± 0.05 ± 0.16		<sup>1</sup> AUBERT	07H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
3.03 ± 0.23 ± 0.23		<sup>2</sup> AUBERT,BE	06J	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
2.68 ± 0.12 ± 0.24		<sup>1,3</sup> AHMED	02B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
2.7 ± 0.6 ± 0.5		<sup>4</sup> BORTOLETTO	92	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
4.8 ± 1.1 ± 1.1	22	<sup>5</sup> ALBRECHT	90J	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
5.1 ± 2.8 ± 1.3 ± 2.5 ± 1.2	4	<sup>6</sup> BEBEK	87	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.79 ± 0.20 ± 0.11		<sup>1,7</sup> AUBERT,B	04o	BABR Repl. by AUBERT 07H
2.8 ± 0.4 ± 0.1	81	<sup>8</sup> ALAM	94	CLE2 Repl. by AHMED 02B
3.1 ± 1.3 ± 1.0	7	<sup>5</sup> ALBRECHT	88k	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> Uses a missing-mass method. Does not depend on $D$ branching fractions or $B^+ / B^0$ production rates.				
<sup>3</sup> AHMED 02b reports an additional uncertainty on the branching ratios to account for 4.5% uncertainty on relative production of $B^0$ and $B^+$ , which is not included here.				
<sup>4</sup> BORTOLETTO 92 assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ and uses Mark III branching fractions for the $D$ .				
<sup>5</sup> ALBRECHT 88k assumes $B^0 \overline{B}^0 : B^+ B^-$ production ratio is 45:55. Superseded by AL-BRECHT 90J which assumes 50:50.				
<sup>6</sup> BEBEK 87 value has been updated in BERKELMAN 91 to use same assumptions as noted for BORTOLETTO 92.				
<sup>7</sup> AUBERT,B 04o reports $[\Gamma(B^0 \rightarrow D^- \pi^+)/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K_S^0 \pi^+)] = (42.7 \pm 2.1 \pm 2.2) \times 10^{-6}$ which we divide by our best value $B(D^+ \rightarrow K_S^0 \pi^+) = (1.53 \pm 0.06) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>8</sup> ALAM 94 reports $[\Gamma(B^0 \rightarrow D^- \pi^+)/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^- 2\pi^+)] = (0.265 \pm 0.032 \pm 0.023) \times 10^{-3}$ which we divide by our best value $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D^- \ell^+ \nu_\ell)/\Gamma(D^- \pi^+)$	$\Gamma_4/\Gamma_{34}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>9.9 ± 1.0 ± 0.9</b>	AALTONEN	09E	CDF	$p\overline{p}$ at 1.96 TeV

$\Gamma(D^- \rho^+)/\Gamma_{\text{total}}$	$\Gamma_{35}/\Gamma$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0075 ± 0.0012 OUR AVERAGE</b>				
0.0074 ± 0.0013 ± 0.0002	79	<sup>1</sup> ALAM	94	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
0.009 ± 0.005 ± 0.003	9	<sup>2</sup> ALBRECHT	90J	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.022 ± 0.012 ± 0.009	6	<sup>2</sup> ALBRECHT	88k	ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALAM 94 reports  $[\Gamma(B^0 \rightarrow D^- \rho^+)/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^- 2\pi^+)] = 0.000704 \pm 0.000096 \pm 0.000070$  which we divide by our best value  $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ALBRECHT 88k assumes  $B^0 \overline{B}^0 : B^+ B^-$  production ratio is 45:55. Superseded by AL-BRECHT 90J which assumes 50:50.



See key on page 601

Meson Particle Listings  
 $B^0$ 

$\Gamma(D^- K^0 \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{36}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>4.9 \pm 0.7 \pm 0.5</math></b>	<sup>1</sup> AUBERT,BE	05B	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^- K^*(892)^+)/\Gamma_{\text{total}}$	$\Gamma_{37}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>4.5 \pm 0.7</math> OUR AVERAGE</b>				
$4.6 \pm 0.6 \pm 0.5$	<sup>1</sup> AUBERT,BE	05B	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$3.7 \pm 1.5 \pm 1.0$	<sup>1</sup> MAHAPATRA	02	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^- \omega \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{38}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.0028 \pm 0.0005 \pm 0.0004</math></b>	<sup>1</sup> ALEXANDER	01B	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The signal is consistent with all observed  $\omega \pi^+$  having proceeded through the  $\rho^+$  resonance at mass  $1349 \pm 25 \pm 5$  MeV and width  $547 \pm 86^{+46}_{-45}$  MeV.

$\Gamma(D^- K^+)/\Gamma_{\text{total}}$	$\Gamma_{39}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.86 \pm 0.20</math> OUR AVERAGE</b>				
$1.89 \pm 0.19 \pm 0.10$	<sup>1</sup> AAIJ	11F	LHCB	$pp$ at 7 TeV
$1.7 \pm 0.4 \pm 0.1$	<sup>2</sup> ABE	01i	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 11F reports  $(2.01 \pm 0.18 \pm 0.14) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^- K^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \pi^+)]$  assuming  $B(B^0 \rightarrow D^- \pi^+) = (2.68 \pm 0.13) \times 10^{-3}$ , which we rescale to our best value  $B(B^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>2</sup> ABE 01i reports  $[\Gamma(B^0 \rightarrow D^- K^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \pi^+)] = (6.8 \pm 1.5 \pm 0.7) \times 10^{-2}$  which we multiply by our best value  $B(B^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D^- K^+)/\Gamma(D^- \pi^+)$	$\Gamma_{39}/\Gamma_{34}$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>8.22 \pm 0.11 \pm 0.25</math></b>	AAIJ	13P	LHCB	$pp$ at 7 TeV

$\Gamma(D^- K^+ \pi^+ \pi^-)/\Gamma(D^- \pi^+ \pi^+ \pi^-)$	$\Gamma_{40}/\Gamma_{46}$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>5.9 \pm 1.1 \pm 0.5</math></b>	AAIJ	12T	LHCB	$pp$ at 7 TeV

$\Gamma(D^- K^+ \bar{K}^0)/\Gamma_{\text{total}}$	$\Gamma_{41}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.1</math></b>	90	<sup>1</sup> DRUTSKOY	02	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^- K^+ \bar{K}^*(892)^0)/\Gamma_{\text{total}}$	$\Gamma_{42}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>8.8 \pm 1.1 \pm 1.5</math></b>	<sup>1</sup> DRUTSKOY	02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{43}/\Gamma$			
VALUE (units $10^{-4}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>8.8 \pm 0.5</math> OUR AVERAGE</b>				
$8.95 \pm 0.15 \pm 0.52$		<sup>1</sup> AAIJ	15Y	LHCB $pp$ at 7, 8 TeV
$8.4 \pm 0.4 \pm 0.8$		<sup>2</sup> KUZMIN	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.0 \pm 0.6 \pm 1.5$		<sup>2,3</sup> SATPATHY	03	BELL Repl. by KUZMIN 07
$< 16$	90	<sup>2</sup> ALAM	94	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$< 70$	90	<sup>4</sup> BORTOLETTO	92	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
$< 340$	90	<sup>5</sup> BEBEK	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
700 ± 500	5	<sup>6</sup> BEHREND	83	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> The second uncertainty combines in quadrature all systematic uncertainties quoted in the paper. AAIJ 15Y reports  $B(B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-) = (8.46 \pm 0.14 \pm 0.49) \times 10^{-4}$  in the kinematic region  $m(\bar{D}^0 \pi^\pm) > 2.1$  GeV which we corrected to the full phase-space dividing by 0.945 from Belle.<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>3</sup> No assumption about the intermediate mechanism is made in the analysis.<sup>4</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ . The product branching fraction into  $D_1^*(2340) \pi$  followed by  $D_1^*(2340) \rightarrow D^0 \pi$  is  $< 0.0001$  at 90% CL and into  $D_2^*(2460) \pi$  followed by  $D_2^*(2460) \rightarrow D^0 \pi$  is  $< 0.0004$  at 90% CL.<sup>5</sup> BEBEK 87 assume the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.  $B(D^0 \rightarrow K^- \pi^+) = (4.2 \pm 0.4 \pm 0.4)\%$  and  $B(D^0 \rightarrow K^- \pi^+ \pi^+ \pi^-) = (9.1 \pm 0.8 \pm 0.8)\%$  were used.<sup>6</sup> Corrected by us using assumptions:  $B(D^0 \rightarrow K^- \pi^+) = (0.042 \pm 0.006)$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = 50\%$ . The product branching ratio is  $B(B^0 \rightarrow \bar{D}^0 \pi^+ \pi^-) B(\bar{D}^0 \rightarrow K^+ \pi^-) = (0.39 \pm 0.26) \times 10^{-2}$ .

$\Gamma(D^*(2010)^- \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{44}/\Gamma$			
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.74 \pm 0.13</math> OUR AVERAGE</b>				

$2.79 \pm 0.08 \pm 0.17$		<sup>1</sup> AUBERT	07H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$2.50 \pm 0.34 \pm 0.13$		<sup>2,3</sup> AUBERT,BE	06J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$2.81 \pm 0.24 \pm 0.05$		<sup>4</sup> BRANDENB...	98	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$2.6 \pm 0.3 \pm 0.4$	82	<sup>5</sup> ALAM	94	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$3.37 \pm 0.96 \pm 0.02$		<sup>6</sup> BORTOLETTO	92	CLEO $e^+e^- \rightarrow \Upsilon(4S)$
$2.36 \pm 0.88 \pm 0.02$	12	<sup>7</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
$2.36^{+1.50}_{-1.10} \pm 0.02$	5	<sup>8</sup> BEBEK	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$10 \pm 4 \pm 1$	8	<sup>9</sup> AKERS	94J	OPAL $e^+e^- \rightarrow Z$
$2.7 \pm 1.4 \pm 1.0$	5	<sup>10</sup> ALBRECHT	87C	ARG $e^+e^- \rightarrow \Upsilon(4S)$
$3.5 \pm 2 \pm 2$		<sup>11</sup> ALBRECHT	86F	ARG $e^+e^- \rightarrow \Upsilon(4S)$
$17 \pm 5 \pm 5$	41	<sup>12</sup> GILES	84	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> AUBERT,BE 06J reports  $[\Gamma(B^0 \rightarrow D^*(2010)^- \pi^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \pi^+)] = 0.99 \pm 0.11 \pm 0.08$  which we multiply by our best value  $B(B^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>3</sup> Uses a missing-mass method. Does not depend on  $D$  branching fractions or  $B^+/B^0$  production rates.<sup>4</sup> BRANDENBURG 98 assume equal production of  $B^+$  and  $B^0$  at  $\Upsilon(4S)$  and use the  $D^*$  reconstruction technique. The first error is their experiment's error and the second error is the systematic error from the PDG 96 value of  $B(D^* \rightarrow D \pi)$ .<sup>5</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II  $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .<sup>6</sup> BORTOLETTO 92 reports  $(4.0 \pm 1.0 \pm 0.7) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- \pi^+)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .<sup>7</sup> ALBRECHT 90J reports  $(2.8 \pm 0.9 \pm 0.6) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- \pi^+)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .<sup>8</sup> BEBEK 87 reports  $(2.8^{+1.5+1.0}_{-1.2-0.6}) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- \pi^+)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Updated in BERKELMAN 91 to use same assumptions as noted for BORTOLETTO 92 and ALBRECHT 90J.<sup>9</sup> Assumes  $B(Z \rightarrow b\bar{b}) = 0.217$  and 38%  $B_d$  production fraction.<sup>10</sup> ALBRECHT 87C use PDG 86 branching ratios for  $D$  and  $D^*(2010)$  and assume  $B(\Upsilon(4S) \rightarrow B^+ B^-) = 55\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = 45\%$ . Superseded by ALBRECHT 90J.<sup>11</sup> ALBRECHT 86F uses pseudomass that is independent of  $D^0$  and  $D^+$  branching ratios.<sup>12</sup> Assumes  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.60^{+0.08}_{-0.15}$ . Assumes  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = 0.40 \pm 0.02$  Does not depend on  $D$  branching ratios.

$\Gamma(D^*(2010)^- \ell^+ \nu_\ell)/\Gamma(D^*(2010)^- \pi^+)$	$\Gamma_6/\Gamma_{44}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>16.5 \pm 2.3 \pm 1.1</math></b>	AALTONEN	09E	CDF	$p\bar{p}$ at 1.96 TeV

$\Gamma(\bar{D}^0 K^+ K^-)/\Gamma(\bar{D}^0 \pi^+ \pi^-)$	$\Gamma_{45}/\Gamma_{43}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.056 \pm 0.011 \pm 0.007</math></b>	AAIJ	12aM	LHCB	$pp$ at 7 TeV

$\Gamma(D^- \pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{46}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.0060 \pm 0.0007</math> OUR FIT</b>				Error includes scale factor of 1.1.
<b><math>0.0080 \pm 0.0021 \pm 0.0014</math></b>	<sup>1</sup> BORTOLETTO	92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(D^- \pi^+ \pi^+ \pi^-)/\Gamma(D^- \pi^+)$	$\Gamma_{46}/\Gamma_{34}$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>2.39 \pm 0.23</math> OUR FIT</b>				
<b><math>2.38 \pm 0.11 \pm 0.21</math></b>	AAIJ	11E	LHCB	$pp$ at 7 TeV

$\Gamma((D^- \pi^+ \pi^+ \pi^-)_{\text{nonresonant}})/\Gamma_{\text{total}}$	$\Gamma_{47}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.0039 \pm 0.0014 \pm 0.0013</math></b>	<sup>1</sup> BORTOLETTO	92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

# Meson Particle Listings

$B^0$

$\Gamma(D^-\pi^+\rho^0)/\Gamma_{\text{total}}$				$\Gamma_{48}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0011±0.0009±0.0004</b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(D^-a_1(1260)^+)/\Gamma_{\text{total}}$				$\Gamma_{49}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0060±0.0022±0.0024</b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .

$\Gamma(D^*(2010)^-\pi^+\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{50}/\Gamma$
VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.0152±0.0052±0.0001</b>	51	<sup>1</sup> ALBRECHT 90j	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
0.015 ±0.008 ±0.008 8 <sup>2</sup> ALBRECHT 87c ARG  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 90j reports  $0.018 \pm 0.004 \pm 0.005$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^-\pi^+\pi^0)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0\pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .  
<sup>2</sup> ALBRECHT 87c use PDG 86 branching ratios for  $D$  and  $D^*(2010)$  and assume  $B(\Upsilon(4S) \rightarrow B^+B^-) = 55\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = 45\%$ . Superseded by ALBRECHT 90j.

$\Gamma(D^*(2010)^-\rho^+)/\Gamma_{\text{total}}$				$\Gamma_{51}/\Gamma$
VALUE (units $10^{-3}$ )	EVTs	DOCUMENT ID	TECN	COMMENT

**2.2 ± 1.8 / 2.7 OUR AVERAGE** Error includes scale factor of 5.2.  
1.48± 0.27± 0.26 / 0.57 <sup>1,2</sup> MATVIENKO 15 BELL  $e^+e^- \rightarrow \Upsilon(4S)$   
6.8 ± 0.3 ± 0.9 <sup>1,3</sup> CSORNA 03 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$   
16.0 ±11.3 ± 0.1 <sup>4</sup> BORTOLETTO92 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$   
5.89± 3.52± 0.04 19 <sup>5</sup> ALBRECHT 90j ARG  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
7.4 ± 1.0 ± 1.4 76 <sup>6,7</sup> ALAM 94 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$   
81 ±29 ±59 / -24 19 <sup>8</sup> CHEN 85 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$  resonance.  
<sup>2</sup> The second uncertainty combines in quadrature the systematic and model uncertainties.  
<sup>3</sup> The second error combines the systematic and theoretical uncertainties in quadrature. CSORNA 03 includes data used in ALAM 94. A full angular fit to three complex helicity amplitudes is performed.  
<sup>4</sup> BORTOLETTO 92 reports  $0.019 \pm 0.008 \pm 0.011$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^-\rho^+)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0\pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .  
<sup>5</sup> ALBRECHT 90j reports  $0.007 \pm 0.003 \pm 0.003$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^-\rho^+)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0\pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .  
<sup>6</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II  $B(D^*(2010)^+ \rightarrow D^0\pi^+)$  and absolute  $B(D^0 \rightarrow K^-\pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^-\pi^+\pi^0)/B(D^0 \rightarrow K^-\pi^+)$  and  $B(D^0 \rightarrow K^-2\pi^+\pi^-)/B(D^0 \rightarrow K^-\pi^+)$ .  
<sup>7</sup> This decay is nearly completely longitudinally polarized,  $\Gamma_L/\Gamma = (93 \pm 5 \pm 5)\%$ , as expected from the factorization hypothesis (ROSNER 90). The nonresonant  $\pi^+\pi^0$  contribution under the  $\rho^+$  is less than 9% at 90% CL.  
<sup>8</sup> Uses  $B(D^* \rightarrow D^0\pi^+) = 0.6 \pm 0.15$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = 0.4$ . Does not depend on  $D$  branching ratios.

$\Gamma(D^*(2010)^-K^+)/\Gamma_{\text{total}}$				$\Gamma_{52}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>2.12±0.15 OUR AVERAGE</b>				

2.13±0.12±0.10 <sup>1</sup> AUBERT 06A BABR  $e^+e^- \rightarrow \Upsilon(4S)$   
2.0 ±0.4 ±0.1 <sup>2</sup> ABE 01i BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 06A reports  $[\Gamma(B^0 \rightarrow D^*(2010)^-K^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^*(2010)^-\pi^+)] = 0.0776 \pm 0.0034 \pm 0.0029$  which we multiply by our best value  $B(B^0 \rightarrow D^*(2010)^-\pi^+) = (2.74 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> ABE 01i reports  $[\Gamma(B^0 \rightarrow D^*(2010)^-K^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^*(2010)^-\pi^+)] = 0.074 \pm 0.015 \pm 0.006$  which we multiply by our best value  $B(B^0 \rightarrow D^*(2010)^-\pi^+) = (2.74 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D^*(2010)^-K^+)/\Gamma(D^*(2010)^-\pi^+)$				$\Gamma_{52}/\Gamma_{44}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>(7.76±0.34±0.26) × 10<sup>-2</sup></b>	AAJ	13A0 LHCb	$pp$ at 7 TeV	

$\Gamma(D^*(2010)^-K^0\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{53}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>3.0±0.7±0.3</b>	<sup>1</sup> AUBERT,BE 05b BABR		$e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^*(2010)^-K^*(892)^+)/\Gamma_{\text{total}}$				$\Gamma_{54}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>3.3±0.6 OUR AVERAGE</b>				

3.2±0.6±0.3 <sup>1</sup> AUBERT,BE 05b BABR  $e^+e^- \rightarrow \Upsilon(4S)$   
3.8±1.3±0.8 <sup>2</sup> MAHAPATRA 02 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and an unpolarized final state.

$\Gamma(D^*(2010)^-K^+\bar{K}^0)/\Gamma_{\text{total}}$				$\Gamma_{55}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.7</b>	90	<sup>1</sup> DRUTSKOY 02 BELL		$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^*(2010)^-K^+\bar{K}^*(892)^0)/\Gamma_{\text{total}}$				$\Gamma_{56}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>12.9±2.2±2.5</b>	<sup>1</sup> DRUTSKOY 02 BELL		$e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

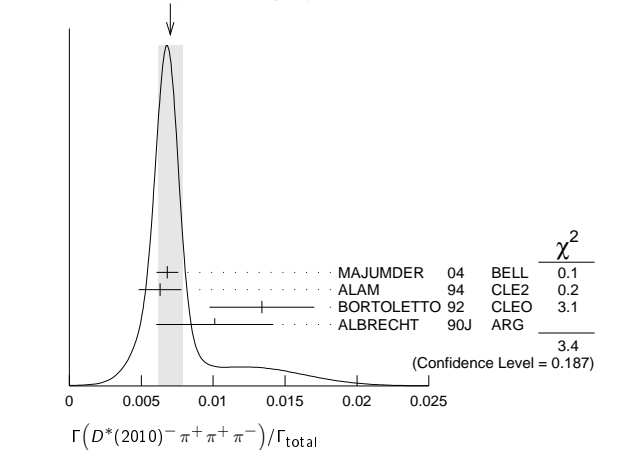
$\Gamma(D^*(2010)^-\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{57}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**0.0070 ±0.0008 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

0.00681 ±0.00023 ±0.00072 <sup>1</sup> MAJUMDER 04 BELL  $e^+e^- \rightarrow \Upsilon(4S)$   
0.0063 ±0.0010 ±0.0011 <sup>2,3</sup> ALAM 94 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$   
0.0134 ±0.0036 ±0.0001 <sup>4</sup> BORTOLETTO92 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$   
0.0101 ±0.0041 ±0.0001 <sup>5</sup> ALBRECHT 90j ARG  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
0.033 ±0.009 ±0.016 <sup>6</sup> ALBRECHT 87c ARG  $e^+e^- \rightarrow \Upsilon(4S)$   
<0.042 90 <sup>7</sup> BEBEK 87 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II  $B(D^*(2010)^+ \rightarrow D^0\pi^+)$  and absolute  $B(D^0 \rightarrow K^-\pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^-\pi^+\pi^0)/B(D^0 \rightarrow K^-\pi^+)$  and  $B(D^0 \rightarrow K^-2\pi^+\pi^-)/B(D^0 \rightarrow K^-\pi^+)$ .  
<sup>3</sup> The three pion mass is required to be between 1.0 and 1.6 GeV consistent with an  $a_1$  meson. (If this channel is dominated by  $a_1^+$ , the branching ratio for  $\bar{D}^*-a_1^+$  is twice that for  $\bar{D}^*-\pi^+\pi^+\pi^-$ .)  
<sup>4</sup> BORTOLETTO 92 reports  $0.0159 \pm 0.0028 \pm 0.0037$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^-\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0\pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .  
<sup>5</sup> ALBRECHT 90j reports  $0.012 \pm 0.003 \pm 0.004$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^-\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0\pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .  
<sup>6</sup> ALBRECHT 87c use PDG 86 branching ratios for  $D$  and  $D^*(2010)$  and assume  $B(\Upsilon(4S) \rightarrow B^+B^-) = 55\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = 45\%$ . Superseded by ALBRECHT 90j.  
<sup>7</sup> BEBEK 87 value has been updated in BERKELMAN 91 to use same assumptions as noted for BORTOLETTO 92.



$\Gamma((D^*(2010)^- \pi^+ \pi^+ \pi^-) \text{ nonresonant})/\Gamma_{\text{total}}$				$\Gamma_{58}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0000<math>\pm</math>0.0019<math>\pm</math>0.0016</b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses MarkIII branching fractions for the  $D$  and  $D^*(2010)$ .

$\Gamma(D^*(2010)^- \pi^+ \rho^0)/\Gamma_{\text{total}}$				$\Gamma_{59}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.00573<math>\pm</math>0.00317<math>\pm</math>0.00004</b>	<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BORTOLETTO 92 reports  $0.0068 \pm 0.0032 \pm 0.0021$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- \pi^+ \rho^0)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses MarkIII branching fractions for the  $D$ .

$\Gamma(D^*(2010)^- a_1(1260)^+)/\Gamma_{\text{total}}$				$\Gamma_{60}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0130<math>\pm</math>0.0027 OUR AVERAGE</b>				
0.0126 $\pm$ 0.0020 $\pm$ 0.0022	<sup>1,2</sup> ALAM	94	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
0.0152 $\pm$ 0.0070 $\pm$ 0.0001	<sup>3</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> ALAM 94 value is twice their  $\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$  value based on their observation that the three pions are dominantly in the  $a_1(1260)$  mass range 1.0 to 1.6 GeV.

<sup>2</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEOII  $B(D^*(2010)^+ \rightarrow D^0 \pi^+)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

<sup>3</sup> BORTOLETTO 92 reports  $0.018 \pm 0.006 \pm 0.006$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- a_1(1260)^+)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses MarkIII branching fractions for the  $D$ .

$\Gamma(\bar{D}_1(2420)^0 \pi^- \pi^+, \bar{D}_1^0 \rightarrow D^{*-} \pi^+)/\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^-)$				$\Gamma_{61}/\Gamma_{57}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>(2.04<math>\pm</math>0.42<math>\pm</math>0.22) <math>\times 10^{-2}</math></b>	AAIJ	13A0	LHCb $pp$ at 7 TeV	

$\Gamma(D^*(2010)^- K^+ \pi^- \pi^+)/\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^-)$				$\Gamma_{62}/\Gamma_{57}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>(6.47<math>\pm</math>0.37<math>\pm</math>0.35) <math>\times 10^{-2}</math></b>	AAIJ	13A0	LHCb $pp$ at 7 TeV	

$\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{63}/\Gamma$
VALUE	EVT5	DOCUMENT ID	TECN	COMMENT
<b>0.0176<math>\pm</math>0.0027 OUR AVERAGE</b>				
0.0172 $\pm$ 0.0014 $\pm$ 0.0024	<sup>1</sup>	ALEXANDER 01B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
0.0345 $\pm$ 0.0181 $\pm$ 0.0003	28	<sup>2</sup> ALBRECHT 90J	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . The signal is consistent with all observed  $\omega \pi^+$  having proceeded through the  $\rho'^+$  resonance at mass  $1349 \pm 25 \pm 10$  MeV and width  $547 \pm 86 \pm 46$  MeV.

<sup>2</sup> ALBRECHT 90J reports  $0.041 \pm 0.015 \pm 0.016$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- \pi^+ \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses MarkIII branching fractions for the  $D$ .

$\Gamma(D^{*-} 3\pi^+ 2\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{64}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>4.72<math>\pm</math>0.59<math>\pm</math>0.71</b>	<sup>1</sup> MAJUMDER 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^*(2010)^- \omega \pi^+)/\Gamma_{\text{total}}$				$\Gamma_{65}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>2.46<math>\pm</math>0.18 OUR AVERAGE</b>	Error includes scale factor of 1.2.			
2.31 $\pm$ 0.11 $\pm$ 0.14	<sup>1</sup> MATVIENKO 15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
2.88 $\pm$ 0.21 $\pm$ 0.31	<sup>1</sup> AUBERT 06L	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
2.9 $\pm$ 0.3 $\pm$ 0.4	<sup>1,2</sup> ALEXANDER 01B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> The signal is consistent with all observed  $\omega \pi^+$  having proceeded through the  $\rho'^+$  resonance at mass  $1349 \pm 25 \pm 10$  MeV and width  $547 \pm 86 \pm 46$  MeV.

$\Gamma(D_1(2430)^0 \omega, D_1^0 \rightarrow D^{*-} \pi^+)/\Gamma_{\text{total}}$				$\Gamma_{66}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	

**2.7 $\pm$ 0.8 $\pm$ 0.4 OUR AVERAGE**

2.5 $\pm$ 0.4 $\pm$ 0.8  
-0.2

<sup>1,2</sup> MATVIENKO 15 BELL  $e^+ e^- \rightarrow \Upsilon(4S)$

4.1 $\pm$ 1.2 $\pm$ 1.1

<sup>3</sup> AUBERT 06L BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$ .

<sup>2</sup> The measurement is obtained by amplitude analysis of  $B^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.

<sup>3</sup> Obtained by fitting the events with  $\cos \theta_{D^*} < 0.5$  and scaling up the result by a factor of 4/3. No interference effects between  $B^0 \rightarrow D_1^+ \omega$  and  $D^* \omega \pi$  are assumed.

$\Gamma(\bar{D}^{*-} \rho(1450)^+)/\Gamma_{\text{total}}$				$\Gamma_{67}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.07<math>\pm</math>0.15<math>\pm</math>0.40 -0.31 -0.13</b>	<sup>1,2</sup> MATVIENKO 15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Obtained by amplitude analysis of  $\bar{B}^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(\bar{D}_1(2420)^0 \omega)/\Gamma_{\text{total}}$				$\Gamma_{68}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>0.7<math>\pm</math>0.2<math>\pm</math>0.1</b>	<sup>1,2</sup> MATVIENKO 15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Obtained by amplitude analysis of  $\bar{B}^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(\bar{D}_2^*(2460)^0 \omega)/\Gamma_{\text{total}}$				$\Gamma_{69}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>0.4<math>\pm</math>0.1<math>\pm</math>0.1</b>	<sup>1,2</sup> MATVIENKO 15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Obtained by amplitude analysis of  $\bar{B}^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^{*-} b_1(1235)^-, b_1^- \rightarrow \omega \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{70}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.7 <math>\times 10^{-4}</math></b>	90	<sup>1</sup> MATVIENKO 15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^{*-} \pi^+)/\Gamma_{\text{total}}$				$\Gamma_{71}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.9<math>\pm</math>0.9<math>\pm</math>0.1</b>	<sup>1,2</sup> AUBERT, BE 06J	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> AUBERT, BE 06J reports  $[\Gamma(B^0 \rightarrow \bar{D}^{*-} \pi^+)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \pi^+)] = 0.77 \pm 0.22 \pm 0.29$  which we multiply by our best value  $B(B^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Uses a missing-mass method. Does not depend on  $D$  branching fractions or  $B^+$ / $B^0$  production rates.

$\Gamma(D_1(2420)^- \pi^+, D_1^- \rightarrow D^- \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{72}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>0.99<math>\pm</math>0.20 -0.25</b> OUR FIT				
<b>0.89<math>\pm</math>0.15<math>\pm</math>0.17 -0.32</b>	<sup>1</sup> ABE 05A	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D_1(2420)^- \pi^+, D_1^- \rightarrow D^- \pi^+ \pi^-)/\Gamma(D^- \pi^+ \pi^- \pi^-)$				$\Gamma_{72}/\Gamma_{46}$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>1.65<math>\pm</math>0.35 -0.40</b> OUR FIT				

2.1  $\pm$  0.5  $\pm$  0.3  
-0.5

AAIJ 11E LHCb  $pp$  at 7 TeV

$\Gamma(D_1(2420)^- \pi^+, D_1^- \rightarrow D^{*-} \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{73}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.33</b>	90	<sup>1</sup> ABE 05A	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^-)/\Gamma(D^*(2010)^- \pi^+)$				$\Gamma_{57}/\Gamma_{44}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>2.64<math>\pm</math>0.04<math>\pm</math>0.13</b>	AAIJ	13A0	LHCb $pp$ at 7 TeV	

$\Gamma(\bar{D}_2^*(2460)^- \pi^+, (D_2^*)^- \rightarrow D^0 \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{74}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.38<math>\pm</math>0.16 OUR AVERAGE</b>				
2.44 $\pm$ 0.07 $\pm$ 0.16	<sup>1</sup> AAJ	15Y	LHCb	$pp$ at 7, 8 TeV
2.15 $\pm$ 0.17 $\pm$ 0.31	<sup>2,3</sup> KUZMIN 07	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<14.7 90 <sup>2</sup> ALAM 94 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Result obtained using the isobar formalism. The second uncertainty combines in quadrature all systematic uncertainties quoted in the paper.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> Our second uncertainty combines systematics and model errors quoted in the paper.

# Meson Particle Listings

$B^0$

$\Gamma(\overline{D}_s^0(2400)^-\pi^+, (D_s^0)^-\rightarrow D^0\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{75}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.76±0.08 OUR AVERAGE</b>				
0.77±0.05±0.06		<sup>1</sup> AAIJ	15Y	LHCB $p\bar{p}$ at 7, 8 TeV
0.60±0.13±0.27		<sup>2,3</sup> KUZMIN	07	BELL $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> Result obtained using the isobar formalism. The second uncertainty combines in quadrature all systematic uncertainties quoted in the paper.
- <sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>3</sup> Our second uncertainty combines systematics and model errors quoted in the paper.

$\Gamma(D_s^{*2}(2460)^-\pi^+, (D_s^{*2})^-\rightarrow D^{*-}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{76}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.24</b>	90	<sup>1</sup> ABE	05A	BELL $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\overline{D}_s^{*2}(2460)^-\rho^+)/\Gamma_{\text{total}}$				$\Gamma_{77}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.0049</b>	90	<sup>1</sup> ALAM	94	CLE2 $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> ALAM 94 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II absolute  $B(D^0\rightarrow K^-\pi^+)$  and  $B(D_s^{*2}(2460)^+\rightarrow D^0\pi^+)=30\%$ .

$\Gamma(D^0\overline{D}^0)/\Gamma_{\text{total}}$				$\Gamma_{78}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.14±0.06±0.03</b>		<sup>1</sup> AAIJ	13AP	LHCB $p\bar{p}$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.43	90	<sup>2</sup> ADACHI	08	BELL $e^+e^-\rightarrow\Upsilon(4S)$
<0.6	90	<sup>2</sup> AUBERT,B	06A	BABR $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> Uses  $B(B^0\rightarrow D^-D^+)=(2.11\pm0.31)\times10^{-4}$  and  $B(B^+\rightarrow\overline{D}^0D_s^+)=(10.1\pm1.7)\times10^{-3}$ .
- <sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^{*0}\overline{D}^0)/\Gamma_{\text{total}}$				$\Gamma_{79}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.9</b>	90	<sup>1</sup> AUBERT,B	06A	BABR $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^-D^+)/\Gamma_{\text{total}}$				$\Gamma_{80}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.11±0.18 OUR AVERAGE</b>				
2.12±0.16±0.18		<sup>1</sup> ROHRKEN	12	BELL $e^+e^-\rightarrow\Upsilon(4S)$
1.97±0.20±0.20		<sup>1</sup> FRATINA	07	BELL $e^+e^-\rightarrow\Upsilon(4S)$
2.8±0.4±0.5		<sup>1</sup> AUBERT,B	06A	BABR $e^+e^-\rightarrow\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.91±0.51±0.30		<sup>1</sup> MAJUMDER	05	BELL Repl. by FRATINA 07
< 9.4	90	<sup>1</sup> LIPLEES	00	CLE2 $e^+e^-\rightarrow\Upsilon(4S)$
<59	90	BARATE	98Q	ALEP $e^+e^-\rightarrow Z$
<12	90	ASNER	97	CLE2 $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^{\pm}D^{*\mp}(CP\text{-averaged}))/\Gamma_{\text{total}}$				$\Gamma_{81}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>6.14±0.29±0.50</b>		<sup>1</sup> ROHRKEN	12	BELL $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D^-D_s^+)/\Gamma_{\text{total}}$				$\Gamma_{82}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0072±0.0008 OUR AVERAGE</b>				
0.0073±0.0004±0.0007		<sup>1</sup> ZUPANC	07	BELL $e^+e^-\rightarrow\Upsilon(4S)$
0.0066±0.0014±0.0006		<sup>2</sup> AUBERT	06N	BABR $e^+e^-\rightarrow\Upsilon(4S)$
0.0068±0.0024±0.0006		<sup>3</sup> GIBAUT	96	CLE2 $e^+e^-\rightarrow\Upsilon(4S)$
0.010±0.009±0.001		<sup>4</sup> ALBRECHT	92G	ARG $e^+e^-\rightarrow\Upsilon(4S)$
0.0053±0.0030±0.0005		<sup>5</sup> BORTOLETTO92	CLEO	$e^+e^-\rightarrow\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.012±0.007	3	<sup>6</sup> BORTOLETTO90	CLEO	$e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> ZUPANC 07 reports  $(7.5\pm0.2\pm1.1)\times10^{-3}$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)= (4.4\pm0.6)\times10^{-2}$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> AUBERT 06N reports  $(0.64\pm0.13\pm0.10)\times10^{-2}$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.0462\pm0.0062$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>3</sup> GIBAUT 96 reports  $0.0087\pm0.0024\pm0.0020$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.035$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

- <sup>4</sup> ALBRECHT 92G reports  $0.017\pm0.013\pm0.006$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.027$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^+$  branching ratios, e.g.,  $B(D^+\rightarrow K^-2\pi^+)=7.7\pm1.0\%$ .
- <sup>5</sup> BORTOLETTO 92 reports  $0.0080\pm0.0045\pm0.0030$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.030\pm0.011$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .
- <sup>6</sup> BORTOLETTO 90 assume  $B(D_s\rightarrow\phi\pi^+)=2\%$ . Superseded by BORTOLETTO 92.

$\Gamma(D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}$				$\Gamma_{83}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0080±0.0011 OUR AVERAGE</b>				
0.0073±0.0013±0.0007		<sup>1</sup> AUBERT	06N	BABR $e^+e^-\rightarrow\Upsilon(4S)$
0.0083±0.0015±0.0007		<sup>2</sup> AUBERT	03I	BABR $e^+e^-\rightarrow\Upsilon(4S)$
0.0088±0.0017±0.0008		<sup>3</sup> AHMED	00B	CLE2 $e^+e^-\rightarrow\Upsilon(4S)$
0.008±0.006±0.001		<sup>4</sup> ALBRECHT	92G	ARG $e^+e^-\rightarrow\Upsilon(4S)$
0.011±0.006±0.001		<sup>5</sup> BORTOLETTO92	CLEO	$e^+e^-\rightarrow\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0072±0.0022±0.0006		<sup>6</sup> GIBAUT	96	CLE2 Repl. by AHMED 00B
0.024±0.014	3	<sup>7</sup> BORTOLETTO90	CLEO	$e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> AUBERT 06N reports  $(0.71\pm0.13\pm0.09)\times10^{-2}$  from a measurement of  $[\Gamma(B^0\rightarrow D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.0462\pm0.0062$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> AUBERT 03I reports  $0.0103\pm0.0014\pm0.0013$  from a measurement of  $[\Gamma(B^0\rightarrow D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.036$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>3</sup> AHMED 00B reports  $0.0110\pm0.0018\pm0.0011$  from a measurement of  $[\Gamma(B^0\rightarrow D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.036$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>4</sup> ALBRECHT 92G reports  $0.014\pm0.010\pm0.003$  from a measurement of  $[\Gamma(B^0\rightarrow D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.027$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^+$  and  $D^{*}(2010)^+$  branching ratios, e.g.,  $B(D^0\rightarrow K^-\pi^+)=3.71\pm0.25\%$ ,  $B(D^+\rightarrow K^-2\pi^+)=7.1\pm1.0\%$ , and  $B(D^{*}(2010)^+\rightarrow D^0\pi^+)=55\pm4\%$ .
- <sup>5</sup> BORTOLETTO 92 reports  $0.016\pm0.009\pm0.006$  from a measurement of  $[\Gamma(B^0\rightarrow D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.030\pm0.011$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$  and  $D^{*}(2010)$ .
- <sup>6</sup> GIBAUT 96 reports  $0.0093\pm0.0023\pm0.0016$  from a measurement of  $[\Gamma(B^0\rightarrow D^{*}(2010)^-D_s^+)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.035$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>7</sup> BORTOLETTO 90 assume  $B(D_s\rightarrow\phi\pi^+)=2\%$ . Superseded by BORTOLETTO 92.

$\Gamma(D^-D_s^{*+})/\Gamma_{\text{total}}$				$\Gamma_{84}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0074±0.0016 OUR AVERAGE</b>				
0.0071±0.0016±0.0006		<sup>1</sup> AUBERT	06N	BABR $e^+e^-\rightarrow\Upsilon(4S)$
0.0078±0.0032±0.0007		<sup>2</sup> GIBAUT	96	CLE2 $e^+e^-\rightarrow\Upsilon(4S)$
0.016±0.012±0.001		<sup>3</sup> ALBRECHT	92G	ARG $e^+e^-\rightarrow\Upsilon(4S)$

- <sup>1</sup> AUBERT 06N reports  $(0.69\pm0.16\pm0.09)\times10^{-2}$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^{*+})/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.0462\pm0.0062$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> GIBAUT 96 reports  $0.0100\pm0.0035\pm0.0022$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^{*+})/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.035$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>3</sup> ALBRECHT 92G reports  $0.027\pm0.017\pm0.009$  from a measurement of  $[\Gamma(B^0\rightarrow D^-D_s^{*+})/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$  assuming  $B(D_s^+\rightarrow\phi\pi^+)=0.027$ , which we rescale to our best value  $B(D_s^+\rightarrow\phi\pi^+)= (4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^+$  branching ratios, e.g.,  $B(D^+\rightarrow K^-2\pi^+)=7.7\pm1.0\%$ .

$\Gamma(D^*(2010)^- D_s^{*+})/\Gamma_{\text{total}}$   $\Gamma_{85}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0177±0.0014 OUR AVERAGE</b>			
0.0173±0.0018±0.0015	<sup>1</sup> AUBERT	06N	BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.0188±0.0009±0.0017	<sup>2</sup> AUBERT	05V	BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.0158±0.0027±0.0014	<sup>3</sup> AUBERT	03I	BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.015 ±0.004 ±0.001	<sup>4</sup> AHMED	00B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
0.016 ±0.009 ±0.001	<sup>5</sup> ALBRECHT	92G	ARG $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.016 ±0.005 ±0.001	<sup>6</sup> GIBAUT	96	CLE2 Repl. by AHMED 00B

<sup>1</sup> AUBERT 06N reports  $(1.68 \pm 0.21 \pm 0.19) \times 10^{-2}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.0462 \pm 0.0062$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> A partial reconstruction technique is used and the result is independent of the particle decay rate of  $D_s^+$  meson. It also provides a model-independent determination of  $B(D_s^+ \rightarrow \phi\pi^+) = (4.81 \pm 0.52 \pm 0.38)\%$ .

<sup>3</sup> AUBERT 03I reports  $0.0197 \pm 0.0015 \pm 0.0030$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> AHMED 00B reports  $0.0182 \pm 0.0037 \pm 0.0025$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>5</sup> ALBRECHT 92G reports  $0.026 \pm 0.014 \pm 0.006$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes PDG 1990  $D^+$  and  $D^*(2010)^+$  branching ratios, e.g.,  $B(D^0 \rightarrow K^-\pi^+) = 3.71 \pm 0.25\%$ ,  $B(D^+ \rightarrow K^-\pi^+) = 7.1 \pm 1.0\%$ , and  $B(D^*(2010)^+ \rightarrow D^0\pi^+) = 55 \pm 4\%$ .

<sup>6</sup> GIBAUT 96 reports  $0.0203 \pm 0.0050 \pm 0.0036$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.035$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $[\Gamma(D^*(2010)^- D_s^+) + \Gamma(D^*(2010)^- D_s^{*+})]/\Gamma_{\text{total}}$   $(\Gamma_{83} + \Gamma_{85})/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.5 ±0.4 OUR AVERAGE</b>				
2.40±0.35±0.22		<sup>1</sup> AUBERT	03I	BABR $e^+e^- \rightarrow \Upsilon(4S)$
3.3 ±0.9 ±0.3	22	<sup>2</sup> BORTOLETTO	90	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 03I reports  $(3.00 \pm 0.19 \pm 0.39) \times 10^{-2}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^+) + \Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})]/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> BORTOLETTO 90 reports  $(7.5 \pm 2.0) \times 10^{-2}$  from a measurement of  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_s^+) + \Gamma(B^0 \rightarrow D^*(2010)^- D_s^{*+})]/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.02$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(D_{s0}(2317)^- K^+, D_{s0}^- \rightarrow D_s^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{86}/\Gamma$ 

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.2±1.4±0.4</b>			
	<sup>1</sup> DRUTSKOY	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> DRUTSKOY 05 reports  $(5.3^{+1.5}_{-1.3} \pm 1.6) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_{s0}(2317)^- K^+, D_{s0}^- \rightarrow D_s^- \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(D_{s0}(2317)^- \pi^+, D_{s0}^- \rightarrow D_s^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{87}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5</b>	90	<sup>1</sup> DRUTSKOY	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_{sJ}(2457)^- K^+, D_{sJ}^- \rightarrow D_s^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{88}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.94</b>	90	<sup>1</sup> DRUTSKOY	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_{sJ}(2457)^- \pi^+, D_{sJ}^- \rightarrow D_s^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{89}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.40</b>	90	<sup>1</sup> DRUTSKOY	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_s^- D_s^+)/\Gamma_{\text{total}}$   $\Gamma_{90}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.6 × 10<sup>-5</sup></b>	90	<sup>1</sup> ZUPANC	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<10 × 10 <sup>-5</sup>	90	<sup>1</sup> AUBERT,BE	05F	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_s^- D_s^+)/\Gamma_{\text{total}}$   $\Gamma_{91}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.3 × 10<sup>-4</sup></b>	90	<sup>1</sup> AUBERT,BE	05F	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_s^- D_s^{*+})/\Gamma_{\text{total}}$   $\Gamma_{92}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.4 × 10<sup>-4</sup></b>	90	<sup>1</sup> AUBERT,BE	05F	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_{s0}^*(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{93}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.04±0.17 OUR AVERAGE</b>			Error includes scale factor of 1.1.

0.97±0.16±0.04 <sup>1,2</sup> CHOI 15A BELL  $e^+e^- \rightarrow \Upsilon(4S)$

1.4 <sup>+0.5</sup><sub>-0.4</sub> ±0.1 <sup>2,3</sup> AUBERT,B 04S BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.69±0.29±0.06 <sup>2,4</sup> KROKOVNY 03B BELL Repl. by CHOI 15A

<sup>1</sup> CHOI 15A reports  $(10.2^{+1.3}_{-1.2} \pm 1.0 \pm 0.4) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_{s0}^*(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow K^+ K^-\pi^+)] \times [B(D^+ \rightarrow K^- 2\pi^+)]$  assuming  $B(D_s^+ \rightarrow K^+ K^-\pi^+) = (5.39 \pm 0.21) \times 10^{-2}$ ,  $B(D^+ \rightarrow K^- 2\pi^+) = (9.13 \pm 0.19) \times 10^{-2}$ , which we rescale to our best values  $B(D_s^+ \rightarrow K^+ K^-\pi^+) = (5.45 \pm 0.17) \times 10^{-2}$ ,  $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> AUBERT,B 04S reports  $(1.8 \pm 0.4^{+0.7}_{-0.5}) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_{s0}^*(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> KROKOVNY 03B reports  $(0.86^{+0.33}_{-0.26} \pm 0.26) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_{s0}^*(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(D_{s0}(2317)^+ D^-, D_{s0}^{*+} \rightarrow D_s^+ \gamma)/\Gamma_{\text{total}}$   $\Gamma_{94}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.95</b>	90	<sup>1</sup> KROKOVNY	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_{s0}(2317)^+ D^*(2010)^-, D_{s0}^{*+} \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{95}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.5±0.4<sup>+0.5</sup><sub>-0.4</sub></b>	<sup>1</sup> AUBERT,B	04S	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

 $\Gamma(D_{sJ}(2457)^+ D^-)/\Gamma_{\text{total}}$   $\Gamma_{96}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.5±1.1 OUR AVERAGE</b>			

2.6±1.5±0.7 <sup>1</sup> AUBERT 06N BABR  $e^+e^- \rightarrow \Upsilon(4S)$

4.8±2.2±1.1 <sup>2,3</sup> AUBERT,B 04S BABR  $e^+e^- \rightarrow \Upsilon(4S)$

3.9±1.5±0.9 <sup>2,4</sup> KROKOVNY 03B BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses a missing-mass method in the events that one of the  $B$  mesons is fully reconstructed.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> AUBERT,B 04S reports  $[\Gamma(B^0 \rightarrow D_{sJ}(2457)^+ D^-)/\Gamma_{\text{total}}] \times [B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0)] = (2.3^{+1.0}_{-0.7} \pm 0.3) \times 10^{-3}$  which we divide by our best value  $B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0) = (48 \pm 11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> KROKOVNY 03B reports  $[\Gamma(B^0 \rightarrow D_{sJ}(2457)^+ D^-)/\Gamma_{\text{total}}] \times [B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0)] = (1.9^{+0.7}_{-0.6} \pm 0.2) \times 10^{-3}$  which we divide by our best value  $B(D_{s1}(2460)^+ \rightarrow D_s^{*+} \pi^0) = (48 \pm 11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

$B^0$

## $\Gamma(D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \gamma)/\Gamma_{\text{total}}$ $\Gamma_{97}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.65^{+0.17}_{-0.14}</math> OUR AVERAGE</b>			
$0.64^{+0.24}_{-0.16} \pm 0.06$	1,2 AUBERT,B	04s	BABR $e^+ e^- \rightarrow \gamma(4S)$
$0.66^{+0.21}_{-0.19} \pm 0.06$	1,3 KROKOVNY	03b	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>2</sup> AUBERT,B 04s reports  $(0.8 \pm 0.2^{+0.3}_{-0.2}) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \gamma)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi \pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi \pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> KROKOVNY 03b reports  $(0.82^{+0.22}_{-0.19} \pm 0.25) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \gamma)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi \pi^+) = 0.036 \pm 0.009$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi \pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \gamma)/\Gamma_{\text{total}}$ $\Gamma_{98}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.60</b>	90	1 KROKOVNY	03b	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{99}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.20</b>	90	1 KROKOVNY	03b	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D_{sJ}(2457)^+ D^-, D_{sJ}^+ \rightarrow D_s^+ \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{100}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.36</b>	90	1 KROKOVNY	03b	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^*(2010)^- D_{sJ}(2457)^+)/\Gamma_{\text{total}}$ $\Gamma_{101}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>9.3 \pm 2.2</math> OUR AVERAGE</b>			
$8.8 \pm 2.0 \pm 1.4$	1 AUBERT	06N	BABR $e^+ e^- \rightarrow \gamma(4S)$
11 $^{+5}_{-4} \pm 3$	2,3 AUBERT,B	04s	BABR $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Uses a missing-mass method in the events that one of the  $B$  mesons is fully reconstructed.  
<sup>2</sup> AUBERT,B 04s reports  $[\Gamma(B^0 \rightarrow D^*(2010)^- D_{sJ}(2457)^+)/\Gamma_{\text{total}}] \times [B(D_{s1}(2460)^+ \rightarrow D_s^+ \pi^0)] = (5.5 \pm 1.2^{+2.2}_{-1.6}) \times 10^{-3}$  which we divide by our best value  $B(D_{s1}(2460)^+ \rightarrow D_s^+ \pi^0) = (48 \pm 11) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D_{sJ}(2457)^+ D^*(2010), D_{sJ}^+ \rightarrow D_s^+ \gamma)/\Gamma_{\text{total}}$ $\Gamma_{102}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.3 \pm 0.3^{+0.9}_{-0.6}</math></b>	1 AUBERT,B	04s	BABR $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $[\Gamma(D^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+) + \Gamma(D^{*+} K^0)]/\Gamma_{\text{total}}$ $\Gamma_{103}/\Gamma = (\Gamma_{104} + \Gamma_{105})/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.75 \pm 0.62 \pm 0.36</math></b>	1,2 AUSHEV	11	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Uses  $\Gamma(D^*(2007)^0 \rightarrow D^0 \pi^0) / \Gamma(D^*(2007)^0 \rightarrow D^0 \gamma) = 1.74 \pm 0.13$  and  $\Gamma(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+) / \Gamma(D_{s1}(2536)^+ \rightarrow D^*(2010)^+ K^0) = 1.36 \pm 0.2$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+)/\Gamma_{\text{total}}$ $\Gamma_{104}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.71 \pm 0.48 \pm 0.32</math></b>		1 AUBERT	08b	BABR $e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5	90	AUBERT	03x	BABR Repl. by AUBERT 08b

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0)/\Gamma_{\text{total}}$ $\Gamma_{105}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.61 \pm 1.03 \pm 0.31</math></b>	1 AUBERT	08b	BABR $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $[\Gamma(D^*(2010)^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+) + \Gamma(D^{*+} K^0)]/\Gamma_{\text{total}}$ $\Gamma_{106}/\Gamma = (\Gamma_{107} + \Gamma_{108})/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.01 \pm 1.21 \pm 0.70</math></b>	1,2 AUSHEV	11	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Uses  $\Gamma(D^*(2007)^0 \rightarrow D^0 \pi^0) / \Gamma(D^*(2007)^0 \rightarrow D^0 \gamma) = 1.74 \pm 0.13$  and  $\Gamma(D_{s1}(2536)^+ \rightarrow D^*(2007)^0 K^+) / \Gamma(D_{s1}(2536)^+ \rightarrow D^*(2010)^+ K^0) = 1.36 \pm 0.2$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^*(2010)^- D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*0} K^+)/\Gamma_{\text{total}}$ $\Gamma_{107}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>3.32 \pm 0.88 \pm 0.66</math></b>		1 AUBERT	08b	BABR $e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<7	90	AUBERT	03x	BABR Repl. by AUBERT 08b

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^{*-} D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0)/\Gamma_{\text{total}}$ $\Gamma_{108}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.00 \pm 1.51 \pm 0.67</math></b>	1 AUBERT	08b	BABR $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^- D_{sJ}(2573)^+, D_{sJ}^+ \rightarrow D^0 K^+)/\Gamma_{\text{total}}$ $\Gamma_{109}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>3.4 \pm 1.7 \pm 0.5</math></b>		1 LEES	15c	BABR $e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<10	90	AUBERT	03x	BABR $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^*(2010)^- D_{sJ}(2573)^+, D_{sJ}^+ \rightarrow D^0 K^+)/\Gamma_{\text{total}}$ $\Gamma_{110}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2</b>	90	AUBERT	03x	BABR $e^+ e^- \rightarrow \gamma(4S)$

## $\Gamma(D^- D_{sJ}(2700)^+, D_{sJ}^+ \rightarrow D^0 K^+)/\Gamma_{\text{total}}$ $\Gamma_{111}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>7.14 \pm 0.96 \pm 0.69</math></b>	1 LEES	15c	BABR $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(D^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{112}/\Gamma$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>7.4 \pm 1.2 \pm 0.4</math></b>	1,2 DAS	10	BELL $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> DAS 10 reports  $[\Gamma(B^0 \rightarrow D^+ \pi^-)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow D^- \pi^+)] = (2.92 \pm 0.38 \pm 0.31) \times 10^{-4}$  which we multiply by our best value  $B(B^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Derived using  $\tan(\theta_C) f_D/f_{D_s} \sqrt{B(B^0 \rightarrow D_s^+ \pi^-)/B(B^0 \rightarrow D^- \pi^+)}$  by assuming the flavor SU(3) symmetry, where  $\theta_C$  is the Cabibbo angle,  $f_D$  ( $f_{D_s}$ ) is the  $D$  ( $D_s$ ) meson decay constant.

## $\Gamma(D_s^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{113}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>21.6 \pm 2.6</math> OUR AVERAGE</b>				
$19.9 \pm 2.6 \pm 1.8$		1 DAS	10	BELL $e^+ e^- \rightarrow \gamma(4S)$
$25 \pm 4 \pm 2$		1 AUBERT	08AJ	BABR $e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$14.0 \pm 3.5 \pm 1.3$		2 AUBERT	07k	BABR Repl. by AUBERT 08AJ
$25 \pm 9 \pm 2$		3 AUBERT	03D	BABR Repl. by AUBERT 07k
$19 \pm 9 \pm 7$		4 KROKOVNY	02	BELL Repl. by DAS 10
< 220	90	5 ALEXANDER	93b	CLE2 $e^+ e^- \rightarrow \gamma(4S)$
<1300	90	6 BORTOLETTO	90	CLEO $e^+ e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>2</sup> AUBERT 07k reports  $[\Gamma(B^0 \rightarrow D_s^+ \pi^-)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)] = (0.63 \pm 0.15 \pm 0.05) \times 10^{-6}$  which we divide by our best value  $B(D_s^+ \rightarrow \phi \pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> AUBERT 03D reports  $[\Gamma(B^0 \rightarrow D_s^+ \pi^-)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)] = (1.13 \pm 0.33 \pm 0.21) \times 10^{-6}$  which we divide by our best value  $B(D_s^+ \rightarrow \phi \pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> KROKOVNY 02 reports  $[\Gamma(B^0 \rightarrow D_s^+ \pi^-)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)] = (0.86^{+0.37}_{-0.30} \pm 0.11) \times 10^{-6}$  which we divide by our best value  $B(D_s^+ \rightarrow \phi \pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>5</sup> ALEXANDER 93b reports  $< 270 \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^+ \pi^-)/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi \pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi \pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi \pi^+) = 4.5 \times 10^{-2}$ .  
<sup>6</sup> BORTOLETTO 90 assume  $B(D_s \rightarrow \phi \pi^+) = 2\%$ .

See key on page 601

Meson Particle Listings  
 $B^0$ 

$\Gamma(D_s^{*+}\pi^-) + \Gamma(D_s^{*-}K^+)/\Gamma_{\text{total}}$		$(\Gamma_{113} + \Gamma_{123})/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
$<1.0 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 93E reports  $< 1.7 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\pi^-) + \Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*+}\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{114}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>2.1 \pm 0.4</math> OUR AVERAGE</b>		Error includes scale factor of 1.4.	
$1.75 \pm 0.34 \pm 0.20$		<sup>1</sup> JOSHI 10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$2.6 \pm 0.5 \pm 0.2$		<sup>1</sup> AUBERT 08AJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.9 \pm 0.7 \pm 0.3$		<sup>2</sup> AUBERT 07K	BABR Repl. by AUBERT 08AJ
$< 4.1$	90	AUBERT 03D	BABR Repl. by AUBERT 07K
$< 40$	90	<sup>3</sup> ALEXANDER 93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AUBERT 07K reports  $[\Gamma(B^0 \rightarrow D_s^{*+}\pi^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)] = (1.32 \pm 0.27 \pm 0.15) \times 10^{-6}$  which we divide by our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> ALEXANDER 93B reports  $< 44 \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\pi^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*+}\pi^-) + \Gamma(D_s^{*-}K^+)/\Gamma_{\text{total}}$		$(\Gamma_{114} + \Gamma_{124})/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
$<7 \times 10^{-4}$	90	<sup>1</sup> ALBRECHT 93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 93E reports  $< 1.2 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\pi^-) + \Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*+}\rho^-)/\Gamma_{\text{total}}$		$\Gamma_{115}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt; 2.4</math></b>	90	<sup>1</sup> AUBERT 08AJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$<130$	90	<sup>2</sup> ALBRECHT 93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$
$< 50$	90	<sup>3</sup> ALEXANDER 93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 93E reports  $< 2.2 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\rho^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .  
<sup>3</sup> ALEXANDER 93B reports  $< 6.6 \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\rho^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*+}\rho^-)/\Gamma_{\text{total}}$		$\Gamma_{116}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>4.1 \pm 1.3 \pm 0.4</math></b>		<sup>1</sup> AUBERT 08AJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<150$	90	<sup>2</sup> ALBRECHT 93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$
$< 60$	90	<sup>3</sup> ALEXANDER 93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 93E reports  $< 2.5 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\rho^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .  
<sup>3</sup> ALEXANDER 93B reports  $< 7.4 \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}\rho^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*+}a_0^-)/\Gamma_{\text{total}}$		$\Gamma_{117}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;1.9</math></b>	90	<sup>1</sup> AUBERT 06X	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D_s^{*+}a_0^-)/\Gamma_{\text{total}}$		$\Gamma_{118}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;3.6</math></b>	90	<sup>1</sup> AUBERT 06X	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(D_s^{*+}a_1(1260)^-)/\Gamma_{\text{total}}$		$\Gamma_{119}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;2.1 \times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT 93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 93E reports  $< 3.5 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*+}a_1(1260)^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*+}a_1(1260)^-)/\Gamma_{\text{total}}$		$\Gamma_{120}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;1.7 \times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT 93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93E reports $< 2.9 \times 10^{-3}$ from a measurement of $[\Gamma(B^0 \rightarrow D_s^{*+}a_1(1260)^-)/\Gamma_{\text{total}}] \times [B(D_s^{*+} \rightarrow \phi\pi^+)]$ assuming $B(D_s^{*+} \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $B(D_s^{*+} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .			

$\Gamma(D_s^{*+}a_2^-)/\Gamma_{\text{total}}$		$\Gamma_{121}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;19</math></b>	90	<sup>1</sup> AUBERT 06X	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(D_s^{*+}a_2^-)/\Gamma_{\text{total}}$		$\Gamma_{122}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;20</math></b>	90	<sup>1</sup> AUBERT 06X	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(D_s^{*-}K^+)/\Gamma_{\text{total}}$		$\Gamma_{123}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>27 \pm 5</math> OUR FIT</b>		Error includes scale factor of 2.7.	
<b><math>22 \pm 5</math> OUR AVERAGE</b>		Error includes scale factor of 1.8.	
$19.1 \pm 2.4 \pm 1.7$		<sup>1</sup> DAS 10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$29 \pm 4 \pm 2$		<sup>1</sup> AUBERT 08AJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$27 \pm 5 \pm 2$		<sup>2</sup> AUBERT 07K	BABR Repl. by AUBERT 08AJ
$26 \pm 10 \pm 2$		<sup>3</sup> AUBERT 03D	BABR Repl. by AUBERT 07K
$36 \pm 11 \pm 3$		<sup>4</sup> KROKOVNY 02	BELL Repl. by DAS 10
$< 190$	90	<sup>5</sup> ALEXANDER 93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$<1300$	90	<sup>6</sup> BORTOLETTO 90	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AUBERT 07K reports  $[\Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*-} \rightarrow \phi\pi^+)] = (1.21 \pm 0.17 \pm 0.11) \times 10^{-6}$  which we divide by our best value  $B(D_s^{*-} \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> AUBERT 03D reports  $[\Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*-} \rightarrow \phi\pi^+)] = (1.16 \pm 0.36 \pm 0.24) \times 10^{-6}$  which we divide by our best value  $B(D_s^{*-} \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> KROKOVNY 02 reports  $[\Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*-} \rightarrow \phi\pi^+)] = (1.61 \pm 0.45 \pm 0.38 \pm 0.21) \times 10^{-6}$  which we divide by our best value  $B(D_s^{*-} \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>5</sup> ALEXANDER 93B reports  $< 230 \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*-} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*-} \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^{*-} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .  
<sup>6</sup> BORTOLETTO 90 assume  $B(D_s^{*-} \rightarrow \phi\pi^+) = 2\%$ .

$\Gamma(D_s^{*-}K^+)/\Gamma_{\text{total}}$		$\Gamma_{124}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>2.19 \pm 0.30</math> OUR AVERAGE</b>			
$2.02 \pm 0.33 \pm 0.22$		<sup>1</sup> JOSHI 10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$2.4 \pm 0.4 \pm 0.2$		<sup>1</sup> AUBERT 08AJ	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.2 \pm 0.6 \pm 0.2$		<sup>2</sup> AUBERT 07K	BABR Repl. by AUBERT 08AJ
$< 2.5$	90	AUBERT 03D	BABR Repl. by AUBERT 07K
$<14$	90	<sup>3</sup> ALEXANDER 93B	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AUBERT 07K reports  $[\Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*-} \rightarrow \phi\pi^+)] = (0.97 \pm 0.24 \pm 0.12) \times 10^{-6}$  which we divide by our best value  $B(D_s^{*-} \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> ALEXANDER 93B reports  $< 17 \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*-}K^+)/\Gamma_{\text{total}}] \times [B(D_s^{*-} \rightarrow \phi\pi^+)]$  assuming  $B(D_s^{*-} \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $B(D_s^{*-} \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*-}K^+)/\Gamma(D_s^{*-}\pi^+)$		$\Gamma_{123}/\Gamma_{34}$	
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN COMMENT
<b><math>1.09 \pm 0.19</math> OUR FIT</b>		Error includes scale factor of 2.6.	
<b><math>1.29 \pm 0.05 \pm 0.08</math></b>		AAIJ 15Ac	LHCB $pp$ at 7, 8 TeV

# Meson Particle Listings

$B^0$

$\Gamma(D_s^- K^*(892)^+)/\Gamma_{\text{total}}$		$\Gamma_{125}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.5 \pm 1.0 \pm 0.4</math></b>		<sup>1</sup> AUBERT	08AJ	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<280	90	<sup>2</sup> ALBRECHT	93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
< 80	90	<sup>3</sup> ALEXANDER	93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 93E reports  $< 4.6 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^- K^*(892)^+)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .  
<sup>3</sup> ALEXANDER 93B reports  $< 9.7 \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^- K^*(892)^+)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*-} K^*(892)^+)/\Gamma_{\text{total}}$		$\Gamma_{126}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.2 \pm 1.4 \pm 0.4</math></b>		<sup>1</sup> AUBERT	08AJ	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<350	90	<sup>2</sup> ALBRECHT	93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
< 90	90	<sup>3</sup> ALEXANDER	93B	CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 93E reports  $< 5.8 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*-} K^*(892)^+)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .  
<sup>3</sup> ALEXANDER 93B reports  $< 11.0 \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*-} K^*(892)^+)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.037$ , which we rescale to our best value  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*-} \pi^+ K^0)/\Gamma_{\text{total}}$		$\Gamma_{127}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.97 \pm 0.14</math> OUR AVERAGE</b>					
$0.94 \pm 0.12 \pm 0.10$		<sup>1</sup> WIECHCZYN...15	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$	
$1.10 \pm 0.26 \pm 0.20$		<sup>1</sup> AUBERT	08G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<40	90	<sup>2</sup> ALBRECHT	93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 93E reports  $< 7.3 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*-} \pi^+ K^0)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^{*-} \pi^+ K^0)/\Gamma_{\text{total}}$		$\Gamma_{128}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 1.10</b>	90	<sup>1</sup> AUBERT	08G	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<25	90	<sup>2</sup> ALBRECHT	93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 93E reports  $< 4.2 \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_s^{*-} \pi^+ K^0)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

$\Gamma(D_s^- K^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{129}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.73 \pm 0.32 \pm 0.35</math></b>		<sup>1</sup> AAIJ	12AX	LHCB	$pp$ at 7 TeV
$1$ AAIJ 12AX reports $[\Gamma(B^0 \rightarrow D_s^- K^+ \pi^+ \pi^-)/\Gamma_{\text{total}}] / [\text{B}(B_s^0 \rightarrow D_s^- K^+ \pi^+ \pi^-)] = 0.54 \pm 0.07 \pm 0.07$ which we multiply by our best value $\text{B}(B_s^0 \rightarrow D_s^- K^+ \pi^+ \pi^-) = (3.2 \pm 0.6) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(D_s^{*-} \pi^+ K^*(892)^0)/\Gamma_{\text{total}}$		$\Gamma_{130}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 3.0 \times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT	93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
$1$ ALBRECHT 93E reports $< 5.0 \times 10^{-3}$ from a measurement of $[\Gamma(B^0 \rightarrow D_s^{*-} \pi^+ K^*(892)^0)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$ assuming $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(D_s^{*-} \pi^+ K^*(892)^0)/\Gamma_{\text{total}}$		$\Gamma_{131}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 1.6 \times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT	93E	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
$1$ ALBRECHT 93E reports $< 2.7 \times 10^{-3}$ from a measurement of $[\Gamma(B^0 \rightarrow D_s^{*-} \pi^+ K^*(892)^0)/\Gamma_{\text{total}}] \times [\text{B}(D_s^+ \rightarrow \phi\pi^+)]$ assuming $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value $\text{B}(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .					

$\Gamma(\bar{D}^0 K^0)/\Gamma_{\text{total}}$		$\Gamma_{132}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>5.2 \pm 0.7</math> OUR AVERAGE</b>					
$5.3 \pm 0.7 \pm 0.3$		<sup>1</sup> AUBERT,B	06L	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$5.0 \pm 1.3 \pm 0.6$		<sup>1</sup> KROKOVNY	03	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$1$ Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^0 K^+ \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{133}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>88 \pm 15 \pm 9</math></b>		<sup>1</sup> AUBERT	06A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1$ Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\bar{D}^0 K^+ \pi^-)/\Gamma(\bar{D}^0 \pi^+ \pi^-)$		$\Gamma_{133}/\Gamma_{43}$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.106 \pm 0.007 \pm 0.008</math></b>		AAIJ	13AQ	LHCB	$pp$ at 7 TeV

$\Gamma(\bar{D}^0 K^*(892)^0)/\Gamma_{\text{total}}$		$\Gamma_{134}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>4.5 \pm 0.6</math> OUR AVERAGE</b>					
$5.4 \pm 0.3 \pm 1.1$		<sup>1,2</sup> AAIJ	15X	LHCB	$pp$ at 7, 8 TeV
$4.0 \pm 0.7 \pm 0.3$		<sup>3</sup> AUBERT,B	06L	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$4.8 \pm 1.1 \pm 0.5$		<sup>3</sup> KROKOVNY	03	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$5.7 \pm 0.9 \pm 0.6$		<sup>3</sup> AUBERT	06A	BABR	Repl. by AUBERT,B 06L

- <sup>1</sup> AAIJ 15X reports  $(5.13 \pm 0.20 \pm 0.15 \pm 0.24 \pm 0.60) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{D}^0 K^*(892)^0)/\Gamma_{\text{total}}] \times [\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-)]$  assuming  $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (9.2 \pm 0.6 \pm 0.7 \pm 0.6) \times 10^{-5}$ , which we rescale to our best value  $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (8.8 \pm 1.7) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Measured via amplitude analysis of  $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via  $D^*(2010)^-$  resonance.  
<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{D}^0 K^*(1410)^0)/\Gamma_{\text{total}}$		$\Gamma_{135}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 6.7 \times 10^{-5}</math></b>	90	<sup>1</sup> AAIJ	15X	LHCB	$pp$ at 7, 8 TeV
$1$ Measured via amplitude analysis of $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via $D^*(2010)^-$ resonance.					

$\Gamma(\bar{D}^0 K_0^*(1430)^0)/\Gamma_{\text{total}}$		$\Gamma_{136}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.7 \pm 0.7 \pm 0.1</math></b>		<sup>1,2</sup> AAIJ	15X	LHCB	$pp$ at 7, 8 TeV
$1$ AAIJ 15X reports $(0.71 \pm 0.27 \pm 0.33 \pm 0.47 \pm 0.08) \times 10^{-5}$ from a measurement of $[\Gamma(B^0 \rightarrow \bar{D}^0 K_0^*(1430)^0)/\Gamma_{\text{total}}] \times [\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-)]$ assuming $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (9.2 \pm 0.6 \pm 0.7 \pm 0.6) \times 10^{-5}$ , which we rescale to our best value $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (8.8 \pm 1.7) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. <sup>2</sup> Measured via amplitude analysis of $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via $D^*(2010)^-$ resonance.					

$\Gamma(\bar{D}^0 K_2^*(1430)^0)/\Gamma_{\text{total}}$		$\Gamma_{137}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.1 \pm 0.8 \pm 0.4</math></b>		<sup>1,2</sup> AAIJ	15X	LHCB	$pp$ at 7, 8 TeV
$1$ AAIJ 15X reports $(2.04 \pm 0.45 \pm 0.30 \pm 0.54 \pm 0.25) \times 10^{-5}$ from a measurement of $[\Gamma(B^0 \rightarrow \bar{D}^0 K_2^*(1430)^0)/\Gamma_{\text{total}}] \times [\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-)]$ assuming $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (9.2 \pm 0.6 \pm 0.7 \pm 0.6) \times 10^{-5}$ , which we rescale to our best value $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (8.8 \pm 1.7) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. <sup>2</sup> Measured via amplitude analysis of $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via $D^*(2010)^-$ resonance.					

$\Gamma(D_0^{*-}(2400)^-, D_0^{*-} \rightarrow \bar{D}^0 \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{138}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.9 \pm 0.8 \pm 0.4</math></b>		<sup>1,2</sup> AAIJ	15X	LHCB	$pp$ at 7, 8 TeV
$1$ AAIJ 15X reports $(1.77 \pm 0.26 \pm 0.19 \pm 0.67 \pm 0.20) \times 10^{-5}$ from a measurement of $[\Gamma(B^0 \rightarrow D_0^{*-}(2400)^-, D_0^{*-} \rightarrow \bar{D}^0 \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-)]$ assuming $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (9.2 \pm 0.6 \pm 0.7 \pm 0.6) \times 10^{-5}$ , which we rescale to our best value $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (8.8 \pm 1.7) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. <sup>2</sup> Measured via amplitude analysis of $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via $D^*(2010)^-$ resonance.					

$\Gamma(D_2^{*-}(2460)^- K^+, D_2^{*-} \rightarrow \bar{D}^0 \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{139}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>20.3 \pm 3.5</math> OUR AVERAGE</b>					
$22 \pm 2 \pm 4$		<sup>1,2</sup> AAIJ	15X	LHCB	$pp$ at 7, 8 TeV
$18.3 \pm 4.0 \pm 3.1$		<sup>3</sup> AUBERT	06A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$



# Meson Particle Listings

## $B^0$

- <sup>1</sup> AAIJ 15x reports  $(2.12 \pm 0.10 \pm 0.11 \pm 0.11 \pm 0.25) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow D_2^{*0}(2460)^- K^+, D_2^{*-} \rightarrow \bar{D}^0 \pi^-)/\Gamma_{\text{total}}] \times [B(B^0 \rightarrow \bar{D}^0 K^+ \pi^-)]$  assuming  $B(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (9.2 \pm 0.6 \pm 0.7 \pm 0.6) \times 10^{-5}$ , which we rescale to our best value  $B(B^0 \rightarrow \bar{D}^0 K^+ \pi^-) = (8.8 \pm 1.7) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> Measured via amplitude analysis of  $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via  $D^*(2010)^-$  resonance.
- <sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(D_2^{*0}(2760)^- K^+, D_2^{*-} \rightarrow \bar{D}^0 \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{140}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.10 \times 10^{-5}</math></b>	90	<sup>1</sup> AAIJ	15x	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Measured via amplitude analysis of $B^0 \rightarrow \bar{D}^0 K^+ \pi^-$ , which excludes contribution from decay via $D^*(2010)^-$ resonance.				

### $\Gamma(\bar{D}^0 K^+ \pi^- \text{ non-resonant})/\Gamma_{\text{total}}$ $\Gamma_{141}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;37</math></b>	90	<sup>1</sup> AUBERT	06A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

### $\Gamma([K^+ K^-]_D K^*(892)^0)/\Gamma(\bar{D}^0 K^*(892)^0)$ $\Gamma_{142}/\Gamma_{134}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.05^{+0.17}_{-0.15} \pm 0.04</math></b>	AAIJ	14BN	LHCB $pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.36^{+0.37}_{-0.32} \pm 0.07$	AAIJ	13L	LHCB Repl. by AAIJ 14BN

### $\Gamma([\pi^+ \pi^-]_D K^*(892)^0)/\Gamma(\bar{D}^0 K^*(892)^0)$ $\Gamma_{143}/\Gamma_{134}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.21^{+0.28}_{-0.25} \pm 0.05</math></b>	AAIJ	14BN	LHCB $pp$ at 7, 8 TeV

### $\Gamma(\bar{D}^0 \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{144}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.63 \pm 0.14</math> OUR AVERAGE</b>				
$2.69 \pm 0.09 \pm 0.13$		<sup>1</sup> LEES	11M	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$2.25 \pm 0.14 \pm 0.35$		<sup>1</sup> BLYTH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$2.74^{+0.36}_{-0.32} \pm 0.55$		<sup>1</sup> COAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.9 \pm 0.2 \pm 0.3$		<sup>1</sup> AUBERT	04B	BABR Repl. by LEES 11M
$3.1 \pm 0.4 \pm 0.5$		<sup>1</sup> ABE	02J	BELL Repl. by BLYTH 06
$<1.2$	90	<sup>2</sup> NEMAT1	98	CLE2 Repl. by COAN 02
$<4.8$	90	<sup>3</sup> ALAM	94	CLE2 Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>2</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.
- <sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

### $\Gamma(\bar{D}^0 \rho^0)/\Gamma_{\text{total}}$ $\Gamma_{145}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>3.21 \pm 0.21</math> OUR AVERAGE</b>				
$3.21 \pm 0.10 \pm 0.21$		<sup>1</sup> AAIJ	15Y	LHCB $pp$ at 7, 8 TeV
$3.19 \pm 0.20 \pm 0.45$		<sup>2,3</sup> KUZMIN	07	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.9 \pm 1.0 \pm 0.4$		<sup>2</sup> SATPATHY	03	BELL Repl. by KUZMIN 07
$< 3.9$	90	<sup>4</sup> NEMAT1	98	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$< 5.5$	90	<sup>5</sup> ALAM	94	CLE2 Repl. by NEMAT1 98
$< 6.0$	90	<sup>6</sup> BORTOLETTO	092	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
$< 27.0$	90	<sup>7</sup> ALBRECHT	88K	ARG $e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Measured using isobar formalism in the decay chain  $B^0 \rightarrow \bar{D}^0 \rho(770)$ ,  $\rho \rightarrow \pi^+ \pi^-$  assuming  $B(\rho(770) \rightarrow \pi^+ \pi^-) = 1$ . The second uncertainty combines in quadrature all systematic uncertainties quoted in the paper.
- <sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>3</sup> Our second uncertainty combines systematics and model errors quoted in the paper.
- <sup>4</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.
- <sup>5</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .
- <sup>6</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses Mark III branching fractions for the  $D$ .
- <sup>7</sup> ALBRECHT 88K reports  $< 0.003$  assuming  $B^0 \bar{B}^0 : B^+ B^-$  production ratio is 45:55. We rescale to 50%.

### $\Gamma(\bar{D}^0 f_2)/\Gamma_{\text{total}}$ $\Gamma_{146}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.56 \pm 0.21</math> OUR AVERAGE</b>				
$1.68 \pm 0.11 \pm 0.21$		<sup>1</sup> AAIJ	15Y	LHCB $pp$ at 7, 8 TeV
$1.20 \pm 0.18 \pm 0.38$		<sup>2,3</sup> KUZMIN	07	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Result obtained using the isobar formalism. The second uncertainty combines in quadrature all systematic uncertainties quoted in the paper. Measured in the decay chain  $B^0 \rightarrow \bar{D}^0 f_2(1270)$ ,  $f_2 \rightarrow \pi^+ \pi^-$ .
- <sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>3</sup> Our second uncertainty combines systematics and model errors quoted in the paper.

### $\Gamma(\bar{D}^0 \eta)/\Gamma_{\text{total}}$ $\Gamma_{147}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.36 \pm 0.32</math> OUR AVERAGE</b>				Error includes scale factor of 2.5.
$2.53 \pm 0.09 \pm 0.11$		<sup>1</sup> LEES	11M	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$1.77 \pm 0.16 \pm 0.21$		<sup>1</sup> BLYTH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.5 \pm 0.2 \pm 0.3$		<sup>1</sup> AUBERT	04B	BABR Repl. by LEES 11M
$1.4^{+0.5}_{-0.4} \pm 0.3$		<sup>1</sup> ABE	02J	BELL Repl. by BLYTH 06
$<1.3$	90	<sup>2</sup> NEMAT1	98	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$<6.8$	90	<sup>3</sup> ALAM	94	CLE2 Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>2</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.
- <sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

### $\Gamma(\bar{D}^0 \eta')/\Gamma_{\text{total}}$ $\Gamma_{148}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.38 \pm 0.16</math> OUR AVERAGE</b>				Error includes scale factor of 1.3.
$1.48 \pm 0.13 \pm 0.07$		<sup>1</sup> LEES	11M	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$1.14 \pm 0.20^{+0.10}_{-0.13}$		<sup>1</sup> SCHUMANN	05	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.7 \pm 0.4 \pm 0.2$		<sup>1</sup> AUBERT	04B	BABR Repl. by LEES 11M
$<9.4$	90	<sup>2</sup> NEMAT1	98	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$<8.6$	90	<sup>3</sup> ALAM	94	CLE2 Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>2</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.
- <sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

### $\Gamma(\bar{D}^0 \eta')/\Gamma(\bar{D}^0 \eta)$ $\Gamma_{148}/\Gamma_{147}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.54 \pm 0.07 \pm 0.01</math></b>	LEES	11M	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.7 \pm 0.2 \pm 0.1$	AUBERT	04B	BABR Repl. by LEES 11M

### $\Gamma(\bar{D}^0 \omega)/\Gamma_{\text{total}}$ $\Gamma_{149}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.54 \pm 0.16</math> OUR AVERAGE</b>				
$2.75 \pm 0.72 \pm 0.35$		<sup>1</sup> AAIJ	15Y	LHCB $pp$ at 7, 8 TeV
$2.57 \pm 0.11 \pm 0.14$		<sup>2</sup> LEES	11M	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$2.37 \pm 0.23 \pm 0.28$		<sup>2</sup> BLYTH	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.0 \pm 0.3 \pm 0.4$		<sup>2</sup> AUBERT	04B	BABR Repl. by LEES 11M
$1.8 \pm 0.5^{+0.4}_{-0.3}$		<sup>2</sup> ABE	02J	BELL Repl. by BLYTH 06
$<5.1$	90	<sup>3</sup> NEMAT1	98	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
$<6.3$	90	<sup>4</sup> ALAM	94	CLE2 Repl. by NEMAT1 98

- <sup>1</sup> Result obtained using the isobar model. The second uncertainty combines in quadrature all systematic uncertainties quoted in the paper.
- <sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- <sup>3</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.
- <sup>4</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and use the CLEO II absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

### $\Gamma(D^0 \phi)/\Gamma_{\text{total}}$ $\Gamma_{150}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;11.6</math></b>	90	<sup>1</sup> AUBERT	07A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(D^0 K^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{151}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;19</math></b>	90	<sup>1</sup> AUBERT	06A	BABR Repl. by AUBERT 09AE
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

# Meson Particle Listings

$B^0$

## $\Gamma(D^0 K^+ \pi^-)/\Gamma(\bar{D}^0 K^+ \pi^-)$ $\Gamma_{151}/\Gamma_{133}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.060±0.034 OUR AVERAGE</b>			
0.045 <sup>+0.056+0.028</sup> <sub>-0.050-0.018</sub>	1,2 NEGISHI	12 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
0.068±0.042	3 AUBERT	09AE BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$

- <sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at the  $\mathcal{T}(4S)$ .  
<sup>2</sup> Uses  $D^0 \rightarrow K^- \pi^+$  mode. Restricts  $K^+ \pi^-$  mass within  $\pm 50$  MeV of the nominal  $K^*0$  mass. Corresponds to the upper limit,  $< 0.16$  at 95% CL.  
<sup>3</sup> Reports a signal at the level of 2.5 standard deviations after combining results from  $D^0 \rightarrow K^+ \pi^-$ ,  $K^+ \pi^- \pi^0$ , and  $K^+ \pi^- \pi^+ \pi^-$ .

## $\Gamma(D^0 K^*(892)^0)/\Gamma_{\text{total}}$ $\Gamma_{152}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1</b>	90	1 AUBERT,B	06L BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.8	90	1 KROKOVNY	03 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .

## $\Gamma(\bar{D}^{*0} \gamma)/\Gamma_{\text{total}}$ $\Gamma_{153}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5 × 10<sup>-5</sup></b>	90	1 AUBERT,B	05Q BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.0 × 10 <sup>-5</sup>	90	1 ARTUSO	00 CLE2	$e^+ e^- \rightarrow \mathcal{T}(4S)$

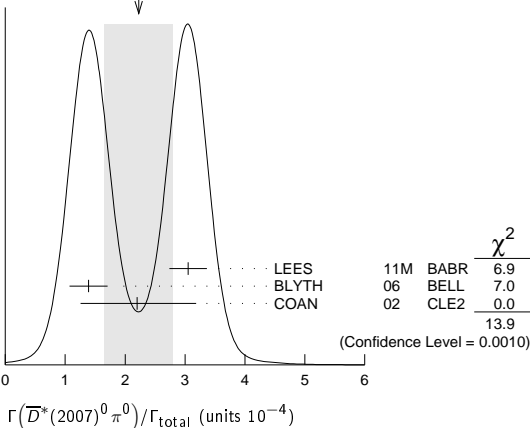
- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .

## $\Gamma(\bar{D}^{*}(2007)^0 \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{154}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.2 ±0.6 OUR AVERAGE</b>				Error includes scale factor of 2.6. See the ideogram below.
3.05±0.14±0.28		1 LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
1.39±0.18±0.26		1 BLYTH	06 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
2.20 <sup>+0.59</sup> <sub>-0.52</sub> ±0.79		1 COAN	02 CLE2	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.9 ±0.4 ±0.5		1 AUBERT	04B BABR	Repl. by LEES 11M
2.7 <sup>+0.8+0.5</sup> <sub>-0.7-0.6</sub>		1 ABE	02J BELL	Repl. by BLYTH 06
<4.4	90	2 NEMAT1	98 CLE2	Repl. by COAN 02
<9.7	90	3 ALAM	94 CLE2	Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .  
<sup>2</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.  
<sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the CLEO11  $B(D^{*}(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

WEIGHTED AVERAGE  
2.2±0.6 (Error scaled by 2.6)



## $\Gamma(\bar{D}^0 \pi^0)/\Gamma(\bar{D}^{*}(2007)^0 \pi^0)$ $\Gamma_{144}/\Gamma_{154}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.90±0.08 OUR AVERAGE</b>			
0.88±0.05±0.06	LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
1.62±0.23±0.35	BLYTH	06 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.0 ±0.1 ±0.2	AUBERT	04B BABR	Repl. by LEES 11M

## $\Gamma(\bar{D}^{*}(2007)^0 \rho^0)/\Gamma_{\text{total}}$ $\Gamma_{155}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.1 × 10<sup>-4</sup></b>	90	1 SATPATHY	03 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.00056	90	2 NEMAT1	98 CLE2	$e^+ e^- \rightarrow \mathcal{T}(4S)$
<0.00117	90	3 ALAM	94 CLE2	Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .  
<sup>2</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.  
<sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the CLEO11  $B(D^{*}(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

## $\Gamma(\bar{D}^{*}(2007)^0 \eta)/\Gamma_{\text{total}}$ $\Gamma_{156}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.3 ±0.6 OUR AVERAGE</b>				Error includes scale factor of 2.8.
2.69±0.14±0.23		1 LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
1.40±0.28±0.26		1 BLYTH	06 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.6 ±0.4 ±0.4		1 AUBERT	04B BABR	Repl. by LEES 11M
<4.6	90	1 ABE	02J BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
<2.6	90	2 NEMAT1	98 CLE2	$e^+ e^- \rightarrow \mathcal{T}(4S)$
<6.9	90	3 ALAM	94 CLE2	Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .  
<sup>2</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.  
<sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the CLEO11  $B(D^{*}(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

## $\Gamma(\bar{D}^0 \eta)/\Gamma(\bar{D}^{*}(2007)^0 \eta)$ $\Gamma_{147}/\Gamma_{156}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.99±0.10 OUR AVERAGE</b>			
0.97±0.07±0.07	LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
1.27±0.29±0.25	BLYTH	06 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.9 ±0.2 ±0.1	AUBERT	04B BABR	Repl. by LEES 11M

## $\Gamma(\bar{D}^{*}(2007)^0 \eta')/\Gamma(\bar{D}^{*}(2007)^0 \eta)$ $\Gamma_{157}/\Gamma_{156}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.61±0.14±0.02</b>	LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.5 ±0.3 ±0.1	AUBERT	04B BABR	Repl. by LEES 11M

## $\Gamma(\bar{D}^{*}(2007)^0 \eta')/\Gamma_{\text{total}}$ $\Gamma_{157}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.40±0.22 OUR AVERAGE</b>				
1.48±0.22±0.13		1 LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
1.21±0.34±0.22		1 SCHUMANN	05 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.3 ±0.7 ±0.2		1,2 AUBERT	04B BABR	Repl. by LEES 11M
<14	90	BRANDENB..	98 CLE2	$e^+ e^- \rightarrow \mathcal{T}(4S)$
<19	90	3 NEMAT1	98 CLE2	$e^+ e^- \rightarrow \mathcal{T}(4S)$
<27	90	4 ALAM	94 CLE2	Repl. by NEMAT1 98

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .  
<sup>2</sup> Reports an upper limit  $< 2.6 \times 10^{-4}$  at 90% CL.  
<sup>3</sup> NEMAT1 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.  
<sup>4</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$  and use the CLEO11  $B(D^{*}(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

## $\Gamma(\bar{D}^0 \eta')/\Gamma(\bar{D}^{*}(2007)^0 \eta')$ $\Gamma_{148}/\Gamma_{157}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.96±0.18±0.06</b>	LEES	11M BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.3 ±0.8 ±0.2	AUBERT	04B BABR	Repl. by LEES 11M

## $\Gamma(\bar{D}^{*}(2007)^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{158}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>(6.2±1.2±1.8) × 10<sup>-4</sup></b>	1,2 SATPATHY	03 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .  
<sup>2</sup> No assumption about the intermediate mechanism is made in the analysis.

## $\Gamma(\bar{D}^{*}(2007)^0 K^0)/\Gamma_{\text{total}}$ $\Gamma_{159}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>3.6±1.2±0.3</b>		1 AUBERT,B	06L BABR	$e^+ e^- \rightarrow \mathcal{T}(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<6.6	90	1 KROKOVNY	03 BELL	$e^+ e^- \rightarrow \mathcal{T}(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\mathcal{T}(4S)$ .

# Meson Particle Listings

 $B^0$ 

$\Gamma(\bar{D}^*(2007)^0 K^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{160}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6.9 \times 10^{-5}$	90	<sup>1</sup> KROKOVNY	03	BELL $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(D^*(2007)^0 K^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{161}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<4.0 \times 10^{-5}$	90	<sup>1</sup> KROKOVNY	03	BELL $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(D^*(2007)^0 \pi^+ \pi^- \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{162}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>2.7 \pm 0.5</math> OUR AVERAGE</b>					
$2.60 \pm 0.47 \pm 0.37$		<sup>1</sup> MAJUMDER	04	BELL $e^+e^- \rightarrow \gamma(4S)$	
$3.0 \pm 0.7 \pm 0.6$		<sup>1</sup> EDWARDS	02	CLE2 $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(D^*(2007)^0 \pi^+ \pi^+ \pi^- \pi^-)/\Gamma(D^*(2010)^- \pi^+ \pi^+ \pi^- \pi^0)$					$\Gamma_{162}/\Gamma_{63}$
VALUE		DOCUMENT ID	TECN	COMMENT	
<b><math>0.17 \pm 0.04 \pm 0.02</math></b>		<sup>1</sup> EDWARDS	02	CLE2 $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(D^*(2010)^+ D^*(2010)^-)/\Gamma_{\text{total}}$					$\Gamma_{163}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>8.0 \pm 0.6</math> OUR AVERAGE</b>					
$7.82 \pm 0.38 \pm 0.63$		<sup>1</sup> KRONENBIT...	12	BELL $e^+e^- \rightarrow \gamma(4S)$	
$8.1 \pm 0.6 \pm 1.0$		<sup>1</sup> AUBERT,B	06A	BABR $e^+e^- \rightarrow \gamma(4S)$	
$9.9 \pm 4.2 \pm 3.3$		<sup>1</sup> LIPELES	00	CLE2 $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$8.1 \pm 0.8 \pm 1.1$		<sup>1</sup> MIYAKE	05	BELL Repl. by KRONENBIT-TER 12	
$8.3 \pm 1.6 \pm 1.2$		<sup>1,2</sup> AUBERT	02M	BABR Repl. by AUBERT,B 06B	
$6.2 \pm 4.0 \pm 2.9$		<sup>3</sup> ARTUSO	99	CLE2 Repl. by LIPELES 00	
$<61$	90	<sup>4</sup> BARATE	98Q	ALEP $e^+e^- \rightarrow Z$	
$<22$	90	<sup>5</sup> ASNER	97	CLE2 Repl. by ARTUSO 99	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .<sup>2</sup> AUBERT 02M also assumes the measured CP-odd fraction of the final states is  $0.22 \pm 0.18 \pm 0.03$ .<sup>3</sup> ARTUSO 99 uses  $B(\gamma(4S) \rightarrow B^0 \bar{B}^0) = (48 \pm 4)\%$ .<sup>4</sup> BARATE 98Q (ALEPH) observes 2 events with an expected background of  $0.10 \pm 0.03$  which corresponds to a branching ratio of  $(2.3 \pm 1.9 \pm 0.4) \times 10^{-3}$ .<sup>5</sup> ASNER 97 at CLEO observes 1 event with an expected background of  $0.022 \pm 0.011$ . This corresponds to a branching ratio of  $(5.3 \pm 7.1 \pm 3.7) \times 10^{-4}$ .

$\Gamma(\bar{D}^*(2007)^0 \omega)/\Gamma_{\text{total}}$					$\Gamma_{164}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.6 \pm 1.1</math> OUR AVERAGE</b>				Error includes scale factor of 3.1.	
$4.55 \pm 0.24 \pm 0.39$		<sup>1</sup> LEES	11M	BABR $e^+e^- \rightarrow \gamma(4S)$	
$2.29 \pm 0.39 \pm 0.40$		<sup>1</sup> BLYTH	06	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$4.2 \pm 0.7 \pm 0.9$	90	<sup>1</sup> AUBERT	04B	BABR Repl. by LEES 11M	
$<7.9$	90	<sup>1</sup> ABE	02J	BELL $e^+e^- \rightarrow \gamma(4S)$	
$<7.4$	90	<sup>2</sup> NEMAT	98	CLE2 $e^+e^- \rightarrow \gamma(4S)$	
$<21$	90	<sup>3</sup> ALAM	94	CLE2 Repl. by NEMAT 98	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .<sup>2</sup> NEMAT 98 assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$  and use the PDG 96 values for  $D^0$ ,  $D^{*0}$ ,  $\eta$ ,  $\eta'$ , and  $\omega$  branching fractions.<sup>3</sup> ALAM 94 assume equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$  and use the CLEO II  $B(D^*(2007)^0 \rightarrow D^0 \pi^0)$  and absolute  $B(D^0 \rightarrow K^- \pi^+)$  and the PDG 1992  $B(D^0 \rightarrow K^- \pi^+ \pi^0)/B(D^0 \rightarrow K^- \pi^+)$  and  $B(D^0 \rightarrow K^- 2\pi^+ \pi^-)/B(D^0 \rightarrow K^- \pi^+)$ .

$\Gamma(\bar{D}^0 \omega)/\Gamma(\bar{D}^*(2007)^0 \omega)$					$\Gamma_{149}/\Gamma_{164}$
VALUE		DOCUMENT ID	TECN	COMMENT	
<b><math>0.58 \pm 0.06</math> OUR AVERAGE</b>					
$0.56 \pm 0.04 \pm 0.04$		LEES	11M	BABR $e^+e^- \rightarrow \gamma(4S)$	
$1.04 \pm 0.20 \pm 0.17$		BLYTH	06	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.7 \pm 0.1 \pm 0.1$		AUBERT	04B	BABR Repl. by LEES 11M	

$\Gamma(D^*(2010)^+ D^-)/\Gamma_{\text{total}}$					$\Gamma_{165}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>6.1 \pm 1.5</math> OUR AVERAGE</b>				Error includes scale factor of 1.6.	
$5.7 \pm 0.7 \pm 0.7$		<sup>1</sup> AUBERT,B	06A	BABR $e^+e^- \rightarrow \gamma(4S)$	
$11.7 \pm 2.6 \pm 2.5$		<sup>1,2</sup> ABE	02Q	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$8.8 \pm 1.0 \pm 1.3$		<sup>1</sup> AUBERT	03J	BABR Repl. by AUBERT,B 06B	
$14.8 \pm 3.8 \pm 3.1$		<sup>1,3</sup> ABE	02Q	BELL $e^+e^- \rightarrow \gamma(4S)$	
$<6.3$	90	<sup>1</sup> LIPELES	00	CLE2 $e^+e^- \rightarrow \gamma(4S)$	
$<56$	90	BARATE	98Q	ALEP $e^+e^- \rightarrow Z$	
$<18$	90	ASNER	97	CLE2 $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .<sup>2</sup> The measurement is performed using fully reconstructed  $D^*$  and  $D^+$  decays.<sup>3</sup> The measurement is performed using a partial reconstruction technique for the  $D^*$  and fully reconstructed  $D^+$  decays as a cross check.

$\Gamma(D^*(2007)^0 \bar{D}^*(2007)^0)/\Gamma_{\text{total}}$					$\Gamma_{166}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.9$	90	<sup>1</sup> AUBERT,B	06A	BABR $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<270$	90	BARATE	98Q	ALEP $e^+e^- \rightarrow Z$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^- D^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{167}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>1.07 \pm 0.07 \pm 0.09</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.7 \pm 0.3 \pm 0.3$		<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^- D^*(2007)^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{168}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>3.46 \pm 0.18 \pm 0.37</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$4.6 \pm 0.7 \pm 0.7$		<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^*(2010)^- D^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{169}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>2.47 \pm 0.10 \pm 0.18</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$3.1 \pm 0.4 \pm 0.3$		<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^*(2010)^- D^*(2007)^0 K^+)/\Gamma_{\text{total}}$					$\Gamma_{170}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>10.6 \pm 0.33 \pm 0.86</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$11.8 \pm 1.0 \pm 1.7$		<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^- D^+ K^0)/\Gamma_{\text{total}}$					$\Gamma_{171}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.75 \pm 0.12 \pm 0.12</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<1.7$	90	<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$[\Gamma(D^*(2010)^- D^+ K^0) + \Gamma(D^- D^*(2010)^+ K^0)]/\Gamma_{\text{total}}$					$\Gamma_{172}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>6.41 \pm 0.36 \pm 0.39</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$6.5 \pm 1.2 \pm 1.0$		<sup>1</sup> AUBERT	03X	BABR Repl. by DEL-AMO-SANCHEZ 11B	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					

$\Gamma(D^*(2010)^- D^*(2010)^+ K^0)/\Gamma_{\text{total}}$					$\Gamma_{173}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>8.1 \pm 0.7</math> OUR AVERAGE</b>					
$8.26 \pm 0.43 \pm 0.67$		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \gamma(4S)$	
$6.8 \pm 0.8 \pm 1.4$		<sup>1,2</sup> DALSENO	07	BELL $e^+e^- \rightarrow \gamma(4S)$	
$8.8 \pm 0.8 \pm 1.4$		<sup>1,2</sup> AUBERT,B	06Q	BABR $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$8.8 \pm 1.5 \pm 1.4$		<sup>1</sup> AUBERT	03X	BABR Repl. by AUBERT,B 06Q	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .<sup>2</sup> The result is rescaled by a factor of 2 to convert from  $K_S^0$  to  $K^0$ .

$\Gamma(D^{*-} D_{s1}(2536)^+, D_{s1}^+ \rightarrow D^{*+} K^0)/\Gamma_{\text{total}}$					$\Gamma_{174}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>8.0 \pm 2.4</math> OUR AVERAGE</b>					
$7.6 \pm 4.8 \pm 1.6$		<sup>1,2</sup> DALSENO	07	BELL $e^+e^- \rightarrow \gamma(4S)$	
$-4.2 \pm 1.4$		<sup>1,2</sup> AUBERT,B	06Q	BABR $e^+e^- \rightarrow \gamma(4S)$	
$8.2 \pm 2.6 \pm 1.2$		<sup>1,2</sup> AUBERT,B	06Q	BABR $e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> The result is rescaled by a factor of 2 to convert from $K_S^0$ to $K^0$ .					

# Meson Particle Listings

$B^0$

$\Gamma(\overline{D}^0 D^0 K^0)/\Gamma_{\text{total}}$					$\Gamma_{175}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.27±0.10±0.05</b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.4	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$[\Gamma(D^0 D^*(2007)^0 K^0) + \Gamma(D^{*-}(2007)^0 D^0 K^0)]/\Gamma_{\text{total}}$					$\Gamma_{176}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.08 \pm 0.32 \pm 0.36</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<3.7	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\overline{D}^*(2007)^0 D^*(2007)^0 K^0)/\Gamma_{\text{total}}$					$\Gamma_{177}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.40 \pm 0.55 \pm 0.67</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<6.6	90	<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma((\overline{D}+\overline{D}^*)(D+D^*)K)/\Gamma_{\text{total}}$					$\Gamma_{178}/\Gamma$
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>3.68 \pm 0.10 \pm 0.24</math></b>		<sup>1</sup> DEL-AMO-SA...11B	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$4.3 \pm 0.3 \pm 0.6$		<sup>1</sup> AUBERT	03x	BABR Repl. by DEL-AMO-SANCHEZ 11B	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\eta_c K^0)/\Gamma_{\text{total}}$					$\Gamma_{179}/\Gamma$	
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT		
<b>0.80±0.12 OUR AVERAGE</b>						
0.55 <sup>+0.19</sup> <sub>-0.18</sub> ± 0.06		<sup>1,2</sup> AUBERT	07AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$		
0.89±0.15 ± 0.06		<sup>1,3</sup> AUBERT,B	04B	BABR $e^+e^- \rightarrow \Upsilon(4S)$		
1.23±0.23 <sup>+0.40</sup> <sub>-0.41</sub>		<sup>1</sup> FANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$		
1.09 <sup>+0.55</sup> <sub>-0.42</sub> ± 0.33		<sup>4</sup> EDWARDS	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$		

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AUBERT 07AV reports  $[\Gamma(B^0 \rightarrow \eta_c K^0)/\Gamma_{\text{total}}] \times [\text{B}(\eta_c(1S) \rightarrow p\overline{p})] = (0.83^{+0.28}_{-0.26} \pm 0.05) \times 10^{-6}$  which we divide by our best value  $\text{B}(\eta_c(1S) \rightarrow p\overline{p}) = (1.50 \pm 0.16) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> AUBERT,B 04B reports  $[\Gamma(B^0 \rightarrow \eta_c K^0)/\Gamma_{\text{total}}] \times [\text{B}(\eta_c(1S) \rightarrow K\overline{K}\pi)] = (0.0648 \pm 0.0085 \pm 0.0071) \times 10^{-3}$  which we divide by our best value  $\text{B}(\eta_c(1S) \rightarrow K\overline{K}\pi) = (7.3 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> EDWARDS 01 assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ . The correlated uncertainties (28.3)% from  $\text{B}(J/\psi(1S) \rightarrow \gamma\eta_c)$  in those modes have been accounted for.

$\Gamma(\eta_c K^0)/\Gamma(J/\psi(1S) K^0)$					$\Gamma_{179}/\Gamma_{183}$	
VALUE		DOCUMENT ID	TECN	COMMENT		
<b>1.39±0.20±0.45</b>		<sup>1</sup> AUBERT,B	04B	BABR $e^+e^- \rightarrow \Upsilon(4S)$		

<sup>1</sup> Uses BABAR measurement of  $\text{B}(B^0 \rightarrow J/\psi K^0) = (8.5 \pm 0.5 \pm 0.6) \times 10^{-4}$ .

$\Gamma(\eta_c K^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{180}/\Gamma$	
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT		
<b>0.63±0.09 OUR AVERAGE</b>						
0.59±0.07 ± 0.07		<sup>1,2</sup> AUBERT	08AB	BABR $e^+e^- \rightarrow \Upsilon(4S)$		
0.69 <sup>+0.21</sup> <sub>-0.20</sub> ± 0.07		<sup>3,4</sup> AUBERT	07AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$		
1.62±0.32 <sup>+0.55</sup> <sub>-0.60</sub>		<sup>4</sup> FANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$		

- <sup>1</sup> AUBERT 08AB reports  $[\Gamma(B^0 \rightarrow \eta_c K^*(892)^0)/\Gamma_{\text{total}}] / [\text{B}(B^+ \rightarrow \eta_c K^+)] = 0.62 \pm 0.06 \pm 0.05$  which we multiply by our best value  $\text{B}(B^+ \rightarrow \eta_c K^+) = (9.6 \pm 1.1) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Uses the production ratio of  $(B^+B^-)/(B^0\overline{B}^0) = 1.026 \pm 0.032$  at  $\Upsilon(4S)$ .  
<sup>3</sup> AUBERT 07AV reports  $[\Gamma(B^0 \rightarrow \eta_c K^*(892)^0)/\Gamma_{\text{total}}] \times [\text{B}(\eta_c(1S) \rightarrow p\overline{p})] = (1.03^{+0.27}_{-0.24} \pm 0.17) \times 10^{-6}$  which we divide by our best value  $\text{B}(\eta_c(1S) \rightarrow p\overline{p}) = (1.50 \pm 0.16) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\eta_c(2S) K^*)/\Gamma_{\text{total}}$					$\Gamma_{181}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;3.9</b>	90	<sup>1</sup> AUBERT	08AB	BABR $e^+e^- \rightarrow \Upsilon(4S)$		

<sup>1</sup> Uses the production ratio of  $(B^+B^-)/(B^0\overline{B}^0) = 1.026 \pm 0.032$  at  $\Upsilon(4S)$ .

$\Gamma(B^0 \rightarrow h_c(1P) K^*)/\Gamma_{\text{total}} \times \Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}}$					$\Gamma_{182}/\Gamma \times \Gamma_{4}^{h_c(1P)}/\Gamma_{h_c(1P)}$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;2.2</b>	90	<sup>1</sup> AUBERT	08AB	BABR $e^+e^- \rightarrow \Upsilon(4S)$		

<sup>1</sup> Uses the production ratio of  $(B^+B^-)/(B^0\overline{B}^0) = 1.026 \pm 0.032$  at  $\Upsilon(4S)$ .

$\Gamma(\eta_c K^*(892)^0)/\Gamma(\eta_c K^0)$					$\Gamma_{180}/\Gamma_{179}$	
VALUE		DOCUMENT ID	TECN	COMMENT		
<b>1.33±0.36<sup>+0.24</sup><sub>-0.33</sub></b>		FANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$		

$\Gamma(J/\psi(1S) K^0)/\Gamma_{\text{total}}$					$\Gamma_{183}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>8.73±0.32 OUR FIT</b>						
<b>8.72±0.32 OUR AVERAGE</b>						
$8.8^{+1.4}_{-1.3} \pm 0.1$			<sup>1,2</sup> AUBERT	07AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$8.69 \pm 0.22 \pm 0.30$			<sup>2</sup> AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
$7.9 \pm 0.4 \pm 0.9$			<sup>2</sup> ABE	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
$9.5 \pm 0.8 \pm 0.6$			<sup>2</sup> AVERY	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
$11.5 \pm 2.3 \pm 1.7$			<sup>3</sup> ABE	96H	CDF $p\bar{p}$ at 1.8 TeV	
$6.93 \pm 4.07 \pm 0.04$			<sup>4</sup> BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	
$9.24 \pm 7.21 \pm 0.05$		<sup>2</sup>	<sup>5</sup> ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$8.3 \pm 0.4 \pm 0.5$			<sup>2</sup> AUBERT	02	BABR Repl. by AUBERT 05J	
$8.5^{+1.4}_{-1.2} \pm 0.6$			<sup>2</sup> JESSOP	97	CLE2 Repl. by AVERY 00	
$7.5 \pm 2.4 \pm 0.8$		<sup>10</sup>	<sup>4</sup> ALAM	94	CLE2 Sup. by JESSOP 97	
<50	90		ALAM	86	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	

- <sup>1</sup> AUBERT 07AV reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^0)/\Gamma_{\text{total}}] \times [\text{B}(J/\psi(1S) \rightarrow p\overline{p})] = (1.87^{+0.28}_{-0.26} \pm 0.07) \times 10^{-6}$  which we divide by our best value  $\text{B}(J/\psi(1S) \rightarrow p\overline{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> ABE 96H assumes that  $\text{B}(B^+ \rightarrow J/\psi K^+) = (1.02 \pm 0.14) \times 10^{-3}$ .  
<sup>4</sup> BORTOLETTO 92 reports  $(6 \pm 3 \pm 2) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^0)/\Gamma_{\text{total}}] \times [\text{B}(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $\text{B}(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $\text{B}(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>5</sup> ALBRECHT 90J reports  $(8 \pm 6 \pm 2) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^0)/\Gamma_{\text{total}}] \times [\text{B}(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $\text{B}(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $\text{B}(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(J/\psi(1S) K^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{184}/\Gamma$	
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>1.15 <math>\pm</math> 0.05 OUR AVERAGE</b>					
1.15 $\pm$ 0.01 $\pm$ 0.05		CHILIKIN 14	BELL	$\overline{B}^0 \rightarrow J/\psi K^- \pi^+$	
1.16 $\pm$ 0.56 $\pm$ 0.01		<sup>1</sup> BORTOLETTO92	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.079 $\pm$ 0.011		<sup>2</sup> AUBERT	09AA	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<1.3	90	<sup>3</sup> ALBRECHT	87D	ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
<6.3	90	GILES	84	CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> BORTOLETTO 92 reports  $(1.0 \pm 0.4 \pm 0.3) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^+\pi^-)/\Gamma_{\text{total}}] \times [\text{B}(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $\text{B}(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $\text{B}(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Does not report systematic uncertainties.  
<sup>3</sup> ALBRECHT 87D assume  $B^+B^-/B^0\overline{B}^0$  ratio is 55/45.  $K\pi$  system is specifically selected as nonresonant.

$\Gamma(J/\psi(1S) K^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{185}/\Gamma$
VALUE (units $10^{-3}$ )		EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.28 ± 0.05</b>	<b>OUR FIT</b>				
<b>1.28 ± 0.05</b>	<b>OUR AVERAGE</b>				
1.19 ± 0.01 ± 0.08			CHILIKIN	14	BELL $\overline{B}^0 \rightarrow J/\psi K^- \pi^+$
$1.33^{+0.22}_{-0.21} \pm 0.02$		<sup>1,2</sup>	AUBERT	07AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.309 ± 0.026 ± 0.077		<sup>2</sup>	AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.29 ± 0.05 ± 0.13		<sup>2</sup>	ABE	02N	BELL $e^+e^- \rightarrow \Upsilon(4S)$
1.74 ± 0.20 ± 0.18		<sup>3</sup>	ABE	98O	CDF $p\overline{p}$ 1.8 TeV
1.32 ± 0.17 ± 0.17		<sup>4</sup>	JESSOP	97	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
1.27 ± 0.65 ± 0.01		<sup>5</sup>	BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
1.27 ± 0.60 ± 0.01		<sup>6</sup>	ALBRECHT	90J	ARG $e^+e^- \rightarrow \Upsilon(4S)$
4.04 ± 1.81 ± 0.02		<sup>5</sup>	BEBEK	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.24 ± 0.05 ± 0.09		<sup>2</sup> AUBERT	02	BABR	Repl. by AUBERT 05J
1.36 ± 0.27 ± 0.22		<sup>8</sup> ABE	96H	CDF	Sup. by ABE 98o
1.69 ± 0.31 ± 0.18	29	<sup>9</sup> ALAM	94	CLE2	Sup. by JESSOP 97
		<sup>10</sup> ALBRECHT	94G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
4.0 ± 0.30		<sup>11</sup> ALBAJAR	91E	UA1	$E_{\text{CM}}^{\text{pp}} = 630 \text{ GeV}$
3.3 ± 0.18	5	<sup>12</sup> ALBRECHT	87D	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
4.1 ± 0.18	5	<sup>13</sup> ALAM	86	CLEO	Repl. by BEBEK 87

<sup>1</sup> AUBERT 07AV reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^*(892)^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow p\bar{p})] = (2.82^{+0.30+0.36}_{-0.28-0.35}) \times 10^{-6}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow p\bar{p}) = (2.120 \pm 0.029) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> ABE 98o reports  $[B(B^0 \rightarrow J/\psi(1S) K^*(892)^0)]/[B(B^+ \rightarrow J/\psi(1S) K^+)] = 1.76 \pm 0.14 \pm 0.15$ . We multiply by our best value  $B(B^+ \rightarrow J/\psi(1S) K^+) = (9.9 \pm 1.0) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>5</sup> BORTOLETTO 92 reports  $(1.1 \pm 0.5 \pm 0.3) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^*(892)^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>6</sup> ALBRECHT 90J reports  $(1.1 \pm 0.5 \pm 0.2) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^*(892)^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>7</sup> BEBEK 87 reports  $(3.5 \pm 1.6 \pm 0.3) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^*(892)^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow e^+e^-)]$  assuming  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.069 \pm 0.009$ , which we rescale to our best value  $B(J/\psi(1S) \rightarrow e^+e^-) = (5.971 \pm 0.032) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Updated in BORTOLETTO 92 to use the same assumptions.

<sup>8</sup> ABE 96H assumes that  $B(B^+ \rightarrow J/\psi K^+) = (1.02 \pm 0.14) \times 10^{-3}$ .

<sup>9</sup> The neutral and charged  $B$  events together are predominantly longitudinally polarized,  $\Gamma_L/\Gamma = 0.080 \pm 0.08 \pm 0.05$ . This can be compared with a prediction using HQET, 0.73 (KRAMER 92). This polarization indicates that the  $B \rightarrow \psi K^*$  decay is dominated by the  $CP = -1$   $CP$  eigenstate. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>10</sup> ALBRECHT 94G measures the polarization in the vector-vector decay to be predominantly longitudinal,  $\Gamma_T/\Gamma = 0.03 \pm 0.16 \pm 0.15$  making the neutral decay a  $CP$  eigenstate when the  $K^*$  decays through  $K_S^0 \pi^0$ .

<sup>11</sup> ALBAJAR 91E assumes  $B_D^0$  production fraction of 36%.

<sup>12</sup> ALBRECHT 87D assume  $B^+B^-/B^0\bar{B}^0$  ratio is 55/45. Superseded by ALBRECHT 90J.

<sup>13</sup> ALAM 86 assumes  $B^\pm/B^0$  ratio is 60/40. The observation of the decay  $B^+ \rightarrow J/\psi K^*(892)^+$  (HAAS 85) has been retracted in this paper.

### $\Gamma(J/\psi(1S) K^*(892)^0)/\Gamma(J/\psi(1S) K^0)$ $\Gamma_{185}/\Gamma_{183}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.50 ± 0.09 OUR AVERAGE</b>			
1.51 ± 0.05 ± 0.08	AUBERT	05J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.39 ± 0.36 ± 0.10	ABE	96Q	CDF $p\bar{p}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.49 ± 0.10 ± 0.08	<sup>1</sup> AUBERT	02	BABR	Repl. by AUBERT 05J
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

### $\Gamma(J/\psi(1S) \eta K_S^0)/\Gamma_{\text{total}}$ $\Gamma_{186}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>5.4 ± 0.9 OUR AVERAGE</b>			
5.22 ± 0.78 ± 0.49	<sup>1</sup> IWASHITA	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
8.4 ± 2.6 ± 2.7	<sup>1</sup> AUBERT	04Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(J/\psi(1S) \eta' K_S^0)/\Gamma_{\text{total}}$ $\Gamma_{187}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.5</b>	90	<sup>1</sup> XIE	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(J/\psi(1S) \omega K^0)/\Gamma_{\text{total}}$ $\Gamma_{189}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.3 ± 0.3 ± 0.3</b>	<sup>1</sup> DEL-AMO-SA...	10B	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.1 ± 0.6 ± 0.3	<sup>1</sup> AUBERT	08W	BABR Repl. by DEL-AMO-SANCHEZ 10B

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(X(3872) K^0, X \rightarrow J/\psi \omega)/\Gamma_{\text{total}}$ $\Gamma_{190}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>6 ± 3 ± 1</b>	<sup>1</sup> DEL-AMO-SA...	10B	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(X(3915), X \rightarrow J/\psi \omega)/\Gamma_{\text{total}}$ $\Gamma_{191}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.1 ± 0.9 ± 0.3</b>	<sup>1</sup> DEL-AMO-SA...	10B	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

1.3 <sup>+1.3</sup> <sub>-1.1</sub> ± 0.2	<sup>1,2</sup> AUBERT	08W	BABR	Repl. by DEL-AMO-SANCHEZ 10B
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> Corresponds to upper limit of $3.9 \times 10^{-5}$ at 90% CL.				

### $\Gamma(J/\psi(1S) \phi K^0)/\Gamma_{\text{total}}$ $\Gamma_{188}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.9 ± 1.0 OUR AVERAGE</b>			Error includes scale factor of 1.3.
4.43 ± 0.76 ± 0.19	LEES	15	BABR $e^+e^- \rightarrow \Upsilon(4S)$
10.2 ± 3.8 ± 1.0	<sup>1</sup> AUBERT	03o	BABR $e^+e^- \rightarrow \Upsilon(4S)$
8.8 <sup>+3.5</sup> <sub>-3.0</sub> ± 1.3	<sup>2</sup> ANASTASSOV	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ANASTASSOV 00 finds 10 events on a background of  $0.5 \pm 0.2$ . Assumes equal production of  $B^0$  and  $B^+$  at the  $\Upsilon(4S)$ , a uniform Dalitz plot distribution, isotropic  $J/\psi(1S)$  and  $\phi$  decays, and  $B(B^+ \rightarrow J/\psi(1S) \phi K^+) = B(B^0 \rightarrow J/\psi(1S) \phi K^0)$ .

### $\Gamma(J/\psi(1S) K(1270)^0)/\Gamma_{\text{total}}$ $\Gamma_{192}/\Gamma$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.30 ± 0.34 ± 0.32</b>	<sup>1</sup> ABE	01L	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and uses the PDG value of  $B(B^+ \rightarrow J/\psi(1S) K^+) = (1.00 \pm 0.10) \times 10^{-3}$ .

### $\Gamma(J/\psi(1S) \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{193}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.76 ± 0.16 OUR AVERAGE</b>				Error includes scale factor of 1.1.
1.69 ± 0.14 ± 0.07		<sup>1</sup> AUBERT	08AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$
2.3 ± 0.5 ± 0.2		<sup>1</sup> ABE	03B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
2.5 <sup>+1.1</sup> <sub>-0.9</sub> ± 0.2		<sup>1</sup> AVERY	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.94 ± 0.22 ± 0.17	<sup>1</sup> AUBERT,B	06B	BABR	Repl. by AUBERT 08AU
2.0 ± 0.6 ± 0.2	<sup>1</sup> AUBERT	02	BABR	Repl. by AUBERT,B 06B
< 32	90	<sup>2</sup> ACCIARRI	97C	L3
< 5.8	90	BISHAI	96	CLE2 Sup. by AVERY 00
< 690	90	<sup>1</sup> ALEXANDER	95	CLE2 Sup. by BISHAI 96

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ACCIARRI 97C assumes  $B_D^0$  production fraction  $(39.5 \pm 4.0\%)$  and  $B_S$   $(12.0 \pm 3.0\%)$ .

### $\Gamma(J/\psi(1S) \eta)/\Gamma_{\text{total}}$ $\Gamma_{194}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>10.8 ± 2.4 OUR AVERAGE</b>				Error includes scale factor of 1.5.
7.3 ± 2.5 ± 1.3		<sup>1</sup> AAIJ	15D	LHCB $pp$ at 7, 8 TeV
12.3 <sup>+1.8</sup> <sub>-1.7</sub> ± 0.7		<sup>2,3</sup> CHANG	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.5 ± 1.7 ± 0.8		<sup>3</sup> CHANG	07A	BELL Repl. by CHANG 12
< 27	90	<sup>3</sup> AUBERT	03o	BABR $e^+e^- \rightarrow \Upsilon(4S)$
< 1200	90	<sup>4</sup> ACCIARRI	97C	L3

<sup>1</sup> AAIJ 15D reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) \eta)/\Gamma_{\text{total}}] / [B(B_S^0 \rightarrow J/\psi(1S) \eta)] = (1.85 \pm 0.61 \pm 0.14) \times 10^{-2}$  which we multiply by our best value  $B(B_S^0 \rightarrow J/\psi(1S) \eta) = (3.9 \pm 0.7) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Reconstructs  $\eta$  in  $\gamma\gamma$  and  $\pi^+\pi^-\pi^0$  decays.

<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>4</sup> ACCIARRI 97C assumes  $B_D^0$  production fraction  $(39.5 \pm 4.0\%)$  and  $B_S$   $(12.0 \pm 3.0\%)$ .

### $\Gamma(J/\psi(1S) \pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{195}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.03 ± 0.18 OUR AVERAGE</b>			
4.00 ± 0.14 ± 0.12	<sup>1,2</sup> AAIJ	13M	LHCB $pp$ at 7 TeV
4.6 ± 0.7 ± 0.6	<sup>3</sup> AUBERT	03B	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 13M reports  $(3.97 \pm 0.09 \pm 0.11 \pm 0.16) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) \pi^+\pi^-)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow J/\psi(1S) K^+)]$  assuming  $B(B^+ \rightarrow J/\psi(1S) K^+) = (1.018 \pm 0.042) \times 10^{-3}$ , which we rescale to our best value  $B(B^+ \rightarrow J/\psi(1S) K^+) = (1.026 \pm 0.031) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> AAIJ 13M does not report correlations between various measurements of the  $J/\psi\pi\pi$  final state.

<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(J/\psi(1S) \pi^+\pi^- \text{ nonresonant})/\Gamma_{\text{total}}$ $\Gamma_{196}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.2</b>	90	<sup>1</sup> AUBERT	07AC	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

# Meson Particle Listings

$B^0$

$\Gamma(J/\psi(1S) f_0(500), f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$				$\Gamma_{197}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	

**$8.1^{+1.1}_{-0.9}$  OUR AVERAGE**

$8.8 \pm 0.5^{+1.1}_{-1.5}$  <sup>1</sup>AAIJ 14x LHCb  $pp$  at 7, 8 TeV

$6.5^{+2.5}_{-1.1} \pm 0.3$  <sup>2,3</sup>AAIJ 13M LHCb  $pp$  at 7 TeV

<sup>1</sup>AAIJ 14x uses Dalitz plot analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

<sup>2</sup>AAIJ 13M reports  $(6.4 \pm 0.8^{+2.4}_{-0.8}) \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) f_0(500), f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-)]$  assuming  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (3.97 \pm 0.09 \pm 0.11 \pm 0.16) \times 10^{-5}$ , which we rescale to our best value  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (4.03 \pm 0.18) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup>AAIJ 13M does not report correlations between various measurements of the  $J/\psi \pi \pi$  final state. Measured in Dalitz plot like analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

$\Gamma(J/\psi(1S) f_2)/\Gamma_{\text{total}}$				$\Gamma_{198}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$0.33^{+0.05}_{-0.06}$  OUR AVERAGE** Error includes scale factor of 1.6.

$0.30 \pm 0.03^{+0.02}_{-0.03}$  <sup>1</sup>AAIJ 14x LHCb  $pp$  at 7, 8 TeV

$0.42 \pm 0.06 \pm 0.02$  <sup>2,3</sup>AAIJ 13M LHCb  $pp$  at 7 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.5 90 <sup>4,5</sup>AUBERT 07AC BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>AAIJ 14x uses Dalitz plot analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

<sup>2</sup>AAIJ 13M reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) f_2)/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] = (3.5 \pm 0.4 \pm 0.4) \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) f_2)/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] / [B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-)]$  assuming  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (3.97 \pm 0.09 \pm 0.11 \pm 0.16) \times 10^{-5}$ , which we rescale to our best values  $B(f_2(1270) \rightarrow \pi\pi) = (84.2^{+2.9}_{-0.9}) \times 10^{-2}$ ,  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (4.03 \pm 0.18) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>3</sup>AAIJ 13M does not report correlations between various measurements of the  $J/\psi \pi \pi$  final state. Measured in Dalitz plot like analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

<sup>4</sup>AUBERT 07AC reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) f_2)/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] < 0.46 \times 10^{-5}$  which we divide by our best value  $B(f_2(1270) \rightarrow \pi\pi) = 84.2 \times 10^{-2}$ .

<sup>5</sup>Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(J/\psi(1S) \rho^0)/\Gamma_{\text{total}}$				$\Gamma_{199}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$2.54 \pm 0.14$  OUR AVERAGE**

$2.50 \pm 0.10^{+0.18}_{-0.15}$  <sup>1</sup>AAIJ 14x LHCb  $pp$  at 7, 8 TeV

$2.52^{+0.22}_{-0.23} \pm 0.11$  <sup>2,3</sup>AAIJ 13M LHCb  $pp$  at 7 TeV

$2.7 \pm 0.3 \pm 0.2$  <sup>4</sup>AUBERT 07AC BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.6 \pm 0.6 \pm 0.4$  <sup>4</sup>AUBERT 03B BABR Repl. by AUBERT 07AC

<25 90 BISHAI 96 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>AAIJ 14x uses Dalitz plot analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ . We assume  $B(\rho(770)^0 \rightarrow \pi^+ \pi^-) = 100\%$ .

<sup>2</sup>AAIJ 13M reports  $(2.49^{+0.20+0.16}_{-0.13-0.23}) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) \rho^0)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-)]$  assuming  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (3.97 \pm 0.09 \pm 0.11 \pm 0.16) \times 10^{-5}$ , which we rescale to our best value  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (4.03 \pm 0.18) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup>AAIJ 13M does not report correlations between various measurements of the  $J/\psi \pi \pi$  final state. Measured in Dalitz plot like analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ . Assumes  $B(\rho(770)^0 \rightarrow \pi\pi) = 100\%$ .

<sup>4</sup>Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(J/\psi(1S) f_0(980), f_0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{200}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT

**$<1.1 \times 10^{-6}$**  <sup>1</sup>AAIJ 90 13M LHCb  $pp$  at 7 TeV

<sup>1</sup>AAIJ 13M does not provide correlations between various measurements of the  $J/\psi \pi^+ \pi^-$  final state. The measurements were obtained from a Dalitz plot like analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ . Also reports  $\Gamma(J/\psi(1S) f_0(980), f_0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}} = (6.1^{+3.1+1.7}_{-2.0-1.4}) \times 10^{-6}$ .

$\Gamma(J/\psi(1S) \rho(1450)^0, \rho^0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$				$\Gamma_{201}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	

**$3.0^{+1.6}_{-0.7}$  OUR AVERAGE**

$4.6 \pm 1.1 \pm 1.9$  <sup>1</sup>AAIJ 14x LHCb  $pp$  at 7, 8 TeV

$2.1^{+2.4}_{-0.7} \pm 0.1$  <sup>2,3</sup>AAIJ 13M LHCb  $pp$  at 7 TeV

<sup>1</sup>AAIJ 14x uses Dalitz plot analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

<sup>2</sup>AAIJ 13M reports  $(2.1^{+1.0+2.2}_{-0.6-0.4}) \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) \rho(1450)^0, \rho^0 \rightarrow \pi\pi)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-)]$  assuming  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (3.97 \pm 0.09 \pm 0.11 \pm 0.16) \times 10^{-5}$ , which we rescale

to our best value  $B(B^0 \rightarrow J/\psi(1S) \pi^+ \pi^-) = (4.03 \pm 0.18) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup>AAIJ 13M does not report correlations between various measurements of the  $J/\psi \pi \pi$  final state. Measured in Dalitz plot like analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

$\Gamma(J/\psi \rho(1700)^0, \rho^0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{202}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	

**$2.0 \pm 0.5 \pm 1.2$**  <sup>1</sup>AAIJ 14x LHCb  $pp$  at 7, 8 TeV

<sup>1</sup>AAIJ 14x uses Dalitz plot analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ .

$\Gamma(J/\psi(1S) \omega)/\Gamma_{\text{total}}$				$\Gamma_{203}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$1.8^{+0.7}_{-0.5} \pm 0.1$**  <sup>1</sup>AAIJ 14x LHCb  $pp$  at 7, 8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<27 90 BISHAI 96 CLE2  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>AAIJ 14x reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) \omega)/\Gamma_{\text{total}}] \times [B(\omega(782) \rightarrow \pi^+ \pi^-)] = (2.7^{+0.8+0.7}_{-0.6-0.5}) \times 10^{-7}$  which we divide by our best value  $B(\omega(782) \rightarrow \pi^+ \pi^-) = (1.53^{+0.11}_{-0.13}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(J/\psi(1S) \omega)/\Gamma(J/\psi(1S) \rho^0)$				$\Gamma_{203}/\Gamma_{199}$
VALUE	DOCUMENT ID	TECN	COMMENT	

**$0.61^{+0.24+0.31}_{-0.14-0.16}$**  <sup>1,2</sup>AAIJ 13M LHCb  $pp$  at 7 TeV

<sup>1</sup>AAIJ 13M reports  $0.61^{+0.24+0.31}_{-0.14-0.16}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) \omega)/\Gamma(B^0 \rightarrow J/\psi(1S) \rho^0)] \times [B(\omega(782) \rightarrow \pi^+ \pi^-)]$  assuming  $B(\omega(782) \rightarrow \pi^+ \pi^-) = (1.53^{+0.11}_{-0.13}) \times 10^{-2}$ .

<sup>2</sup>AAIJ 13M does not report correlations between various measurements of the  $J/\psi \pi \pi$  final state. Measured in Dalitz plot like analysis of  $B^0 \rightarrow J/\psi \pi^+ \pi^-$ . Assumes  $B(\rho(770)^0 \rightarrow \pi\pi) = 100\%$ .

$\Gamma(J/\psi(1S) \omega)/\Gamma(J/\psi(1S) \rho^0)$				$\Gamma_{203}/\Gamma_{199}$
VALUE	DOCUMENT ID	TECN	COMMENT	

**$0.89 \pm 0.19^{+0.07}_{-0.13}$**  AAIJ 13A LHCb  $pp$  at 7 TeV

$\Gamma(J/\psi(1S) K^+ K^-)/\Gamma_{\text{total}}$				$\Gamma_{204}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	

**$2.55 \pm 0.35 \pm 0.08$**  <sup>1</sup>AAIJ 13BT LHCb  $pp$  at 7 TeV

<sup>1</sup>AAIJ 13BT reports  $(2.53 \pm 0.31 \pm 0.19) \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S) K^+ K^-)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow J/\psi(1S) K^+)]$  assuming  $B(B^+ \rightarrow J/\psi(1S) K^+) = (1.018 \pm 0.042) \times 10^{-3}$ , which we rescale to our best value  $B(B^+ \rightarrow J/\psi(1S) K^+) = (1.026 \pm 0.031) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(J/\psi(1S) a_0(980), a_0 \rightarrow K^+ K^-)/\Gamma_{\text{total}}$				$\Gamma_{205}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	

**$0.470 \pm 0.331 \pm 0.072$**  <sup>1</sup>AAIJ 13BT LHCb  $pp$  at 7 TeV

<sup>1</sup>AAIJ 13BT uses  $B(\overline{B}^0 \rightarrow J/\psi K^+ K^-) = (2.53 \pm 0.31 \pm 0.19) \times 10^{-6}$  to derive this result. It also reports the equivalent upper limit of  $< 9.0 \times 10^{-7}$  at 90% CL.

$\Gamma(J/\psi(1S) \phi)/\Gamma_{\text{total}}$				$\Gamma_{206}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**<0.19** 90 <sup>1</sup>AAIJ 13BT LHCb  $pp$  at 7 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.01 90 LEES 15 BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<0.94 90 <sup>2</sup>LIU 08i BELL  $e^+ e^- \rightarrow \Upsilon(4S)$

<9.2 90 <sup>2</sup>AUBERT 03o BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>AAIJ 13BT uses  $B(B^0 \rightarrow J/\psi(1S) K^+ K^-) = (2.53 \pm 0.31 \pm 0.19) \times 10^{-6}$  and  $B(\phi \rightarrow K^+ K^-) = (48.9 \pm 0.5)\%$  to obtain this result.

<sup>2</sup>Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(J/\psi(1S) \eta'(958))/\Gamma_{\text{total}}$				$\Gamma_{207}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT

**$7.6 \pm 2.2^{+0.9}_{-1.0}$**  <sup>1</sup>AAIJ 15D LHCb  $pp$  at 7, 8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 7.4 90 <sup>2,3</sup>CHANG 12 BELL  $e^+ e^- \rightarrow \Upsilon(4S)$

<63 90 <sup>3</sup>AUBERT 03o BABR  $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>AAIJ 15D reports  $[\Gamma(B^0 \rightarrow J/\psi(1S) \eta'(958))/\Gamma_{\text{total}}] / [B(B_S^0 \rightarrow J/\psi(1S) \eta')] = (2.28 \pm 0.65 \pm 0.16) \times 10^{-2}$  which we multiply by our best value  $B(B_S^0 \rightarrow J/\psi(1S) \eta') = (3.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup>Reconstructs  $\eta'(985)$  in  $\eta \pi^+ \pi^-$  and  $\rho(770)^0 \gamma$  decays.

<sup>3</sup>Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

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## Meson Particle Listings

 $B^0$ 

$\Gamma(J/\psi(1S)\eta)/\Gamma(J/\psi(1S)\eta'(958))$				$\Gamma_{194}/\Gamma_{207}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.111 \pm 0.475 \pm 0.062</math></b>	<sup>1</sup> AAIJ	15D	LHCB $pp$ at 7, 8 TeV	

<sup>1</sup> Uses  $J/\psi \rightarrow \mu^+ \mu^-, \eta' \rightarrow \rho^0 \gamma$ , and  $\eta' \rightarrow \eta \pi^+ \pi^-$  decays.

$\Gamma(J/\psi(1S)K^0 \pi^+ \pi^-)/\Gamma(J/\psi(1S)K^0)$				$\Gamma_{208}/\Gamma_{183}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.50 \pm 0.04</math> OUR AVERAGE</b>				
$0.493 \pm 0.034 \pm 0.027$	AAIJ	14L	LHCB $pp$ at 7 TeV	
$0.24 \pm 0.40 \pm 0.15$	AFFOLDER	02B	CDF $p\bar{p}$ 1.8 TeV	

$\Gamma(J/\psi(1S)K^0 K^+ K^-)/\Gamma_{\text{total}}$				$\Gamma_{210}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>25 \pm 7</math> OUR AVERAGE</b>			Error includes scale factor of 1.8.	
$34.9 \pm 6.7 \pm 1.5$	LEES	15	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
$20.2 \pm 4.3 \pm 1.9$	<sup>1</sup> AAIJ	14L	LHCB $pp$ at 7 TeV	

<sup>1</sup> Measured with  $B(B^0 \rightarrow J/\psi K_S^0 K^+ K^-) / B(B^0 \rightarrow J/\psi K_S^0)$  using PDG 12 for the involved branching fractions.

$\Gamma(J/\psi(1S)K^0 K^- \pi^+ + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_{209}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 21 \times 10^{-6}</math></b>	90	<sup>1</sup> AAIJ	14L	LHCB $pp$ at 7 TeV

<sup>1</sup> Measured with  $B(B^0 \rightarrow J/\psi K_S^0 K^\pm \pi^\mp) / B(B^0 \rightarrow J/\psi K_S^0 \pi^+ \pi^-)$  using PDG 12 values for the involved branching fractions.

$\Gamma(J/\psi(1S)K^0 \rho^0)/\Gamma_{\text{total}}$				$\Gamma_{212}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>5.4 \pm 2.9 \pm 0.9</math></b>	<sup>1</sup> AFFOLDER	02B	CDF $p\bar{p}$ 1.8 TeV	

<sup>1</sup> Uses  $B^0 \rightarrow J/\psi(1S)K_S^0$  decay as a reference and  $B(B^0 \rightarrow J/\psi(1S)K^0) = 8.3 \times 10^{-4}$ .

$\Gamma(J/\psi(1S)K^*(892)^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{213}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>7.7 \pm 4.1 \pm 1.3</math></b>	<sup>1</sup> AFFOLDER	02B	CDF $p\bar{p}$ 1.8 TeV	

<sup>1</sup> Uses  $B^0 \rightarrow J/\psi(1S)K_S^0$  decay as a reference and  $B(B^0 \rightarrow J/\psi(1S)K^0) = 8.3 \times 10^{-4}$ .

$\Gamma(J/\psi(1S)\pi^+ \pi^- \pi^+ \pi^-)/\Gamma(J/\psi(1S)\pi^+ \pi^-)$				$\Gamma_{214}/\Gamma_{195}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.361 \pm 0.017 \pm 0.021</math></b>	<sup>1</sup> AAIJ	14Y	LHCB $pp$ at 7, 8 TeV	

<sup>1</sup> Excludes contributions from  $\psi(2S)$  and  $X(3872)$  decaying to  $J/\psi(1S)\pi^+ \pi^-$ .

$\Gamma(J/\psi(1S)f_1(1285))/\Gamma_{\text{total}}$				$\Gamma_{215}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>8.4 \pm 1.9^{+0.8}_{-0.7}</math></b>	<sup>1</sup> AAIJ	14Y	LHCB $pp$ at 7, 8 TeV	

<sup>1</sup> AAIJ 14Y reports  $(8.37 \pm 1.95 \pm 0.71^{+0.71}_{-0.66} \pm 0.35) \times 10^{-6}$  from a measurement of  $[\Gamma(B^0 \rightarrow J/\psi(1S)f_1(1285))/\Gamma_{\text{total}}] \times [B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-)]$  assuming  $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.11 \pm 0.007^{+0.007}_{-0.006}$ .

$\Gamma(J/\psi(1S)K^*(892)^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{216}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>6.6 \pm 1.9 \pm 1.1</math></b>	<sup>1</sup> AFFOLDER	02B	CDF $p\bar{p}$ 1.8 TeV	

<sup>1</sup> Uses  $B^0 \rightarrow J/\psi(1S)K^*(892)^0$  decay as a reference and  $B(B^0 \rightarrow J/\psi(1S)K^0) = 12.4 \times 10^{-4}$ .

$\Gamma(X(3872)^- K^+)/\Gamma_{\text{total}}$				$\Gamma_{217}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 5 \times 10^{-4}</math></b>	90	<sup>1</sup> AUBERT	06E	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Perform measurements of absolute branching fractions using a missing mass technique.

$\Gamma(X(3872)^- K^+, X(3872)^- \rightarrow J/\psi(1S)\pi^- \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{218}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.2</math></b>	90	<sup>1,2</sup> CHOI	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 5.4$	90	<sup>2,3</sup> AUBERT	05B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes  $\pi^+ \pi^0$  originates from  $\rho^+$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> The isovector- $X$  hypothesis is excluded with a likelihood test at  $1 \times 10^{-4}$  level.

$\Gamma(X(3872)K^0, X \rightarrow J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{219}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.3 \pm 1.2 \pm 0.4</math></b>		<sup>1,2</sup> CHOI	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 6.0$	90	<sup>2</sup> AUBERT	08Y	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$< 10.3$	90	<sup>2,3</sup> AUBERT	06	BABR Repl. by AUBERT 08Y

<sup>1</sup> CHOI 11 reports  $[\Gamma(B^0 \rightarrow X(3872)K^0, X \rightarrow J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}] / [B(B^+ \rightarrow X(3872)K^+, X \rightarrow J/\psi \pi^+ \pi^-)] = 0.50 \pm 0.14 \pm 0.04$  which we multiply by our best value  $B(B^+ \rightarrow X(3872)K^+, X \rightarrow J/\psi \pi^+ \pi^-) = (8.6 \pm 0.8) \times 10^{-6}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> The lower limit is also given to be  $1.34 \times 10^{-6}$  at 90% CL.

$\Gamma(X(3872)K^0, X \rightarrow J/\psi \gamma)/\Gamma_{\text{total}}$				$\Gamma_{220}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.4</math></b>	90	<sup>1</sup> BHARDWAJ	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 4.9$	90	<sup>2</sup> AUBERT	09B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872)K^*(892)^0, X \rightarrow J/\psi \gamma)/\Gamma_{\text{total}}$				$\Gamma_{221}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.8</math></b>	90	<sup>1</sup> AUBERT	09B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872)K^0, X \rightarrow \psi(2S)\gamma)/\Gamma_{\text{total}}$				$\Gamma_{222}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 6.62</math></b>	90	<sup>1</sup> BHARDWAJ	11	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 19$	90	<sup>2</sup> AUBERT	09B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872)K^*(892)^0, X \rightarrow \psi(2S)\gamma)/\Gamma_{\text{total}}$				$\Gamma_{223}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.4</math></b>	90	<sup>1</sup> AUBERT	09B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+ B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0 \bar{B}^0) = (48.4 \pm 0.6)\%$ .

$\Gamma(X(3872)K^0, X \rightarrow D^0 \bar{D}^0 \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{224}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.66 \pm 0.70^{+0.32}_{-0.37}</math></b>	<sup>1</sup> GOKHROO	06	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Measure the near-threshold enhancements in the  $(D^0 \bar{D}^0 \pi^0)$  system at a mass  $3875.2 \pm 0.7^{+0.3}_{-1.6} \pm 0.8$  MeV/c<sup>2</sup>.

$\Gamma(X(3872)K^0, X \rightarrow \bar{D}^{*0} D^0)/\Gamma_{\text{total}}$				$\Gamma_{225}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>1.2 \pm 0.4</math> OUR AVERAGE</b>				
$0.97 \pm 0.46 \pm 0.13$	<sup>1</sup> AUSHEV	10	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
$2.22 \pm 1.05 \pm 0.42$	<sup>1,2</sup> AUBERT	08B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> This result is equivalent to the the 90% CL upper limit of  $4.37 \times 10^{-4}$ .

$\Gamma(X(3872)K^+ \pi^-, X \rightarrow J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{226}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>7.9 \pm 1.3 \pm 0.4</math></b>	<sup>1</sup> BALA	15	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3872)K^*(982)^0, X \rightarrow J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_{227}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>4.0 \pm 1.5 \pm 0.3</math></b>	BALA	15	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	

$\Gamma(X(4430)^\pm K^\mp, X^\pm \rightarrow \psi(2S)\pi^\pm)/\Gamma_{\text{total}}$				$\Gamma_{228}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>6.0^{+1.7+2.5}_{-2.0-1.4}</math></b>		CHILIKIN	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 3.1$	95	<sup>1</sup> AUBERT	09AA	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$3.2^{+1.8+5.3}_{-0.9-1.6}$		<sup>1</sup> MIZUK	09	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$4.1 \pm 1.0 \pm 1.4$		<sup>1,2</sup> CHOI	08	BELL Repl. by MIZUK 09

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Establishes the  $X(4430)^+$  with a significance of 6.5 sigma. Needs confirmation.

$\Gamma(X(4430)^\pm K^\mp, X^\pm \rightarrow J/\psi \pi^\pm)/\Gamma_{\text{total}}$				$\Gamma_{229}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.4^{+4.0+1.1}_{-1.0-0.6}</math></b>		CHILIKIN	14	BELL $\bar{B}^0 \rightarrow J/\psi K^- \pi^+$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 4$	95	<sup>1</sup> AUBERT	09AA	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(X(3900)^\pm K^\mp, X^\pm \rightarrow J/\psi \pi^\pm)/\Gamma_{\text{total}}$				$\Gamma_{230}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 9 \times 10^{-7}</math></b>	CHILIKIN	14	BELL $\bar{B}^0 \rightarrow J/\psi K^- \pi^+$	

$\Gamma(X(4200)^\pm K^\mp, X^\pm \rightarrow J/\psi \pi^\pm)/\Gamma_{\text{total}}$				$\Gamma_{231}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>2.2^{+0.7+1.1}_{-0.5-0.6}</math></b>	CHILIKIN	14	BELL $\bar{B}^0 \rightarrow J/\psi K^- \pi^+$	





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# Meson Particle Listings

$B^0$

$\Gamma(\chi_{c2} K^*(892)^0)/\Gamma_{\text{total}}$	$\Gamma_{245}/\Gamma$
VALUE (units $10^{-5}$ )	CL%

**4.9±1.2 OUR FIT** Error includes scale factor of 1.1.  
**6.6±1.8±0.5** <sup>1</sup>AUBERT 09B BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<7.1 90 <sup>2</sup>SONI 06 BELL  $e^+e^- \rightarrow \Upsilon(4S)$   
 <3.6 90 <sup>2</sup>AUBERT 05K BABR Repl. by AUBERT 09B

<sup>1</sup> Uses  $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c2} K^*(892)^0)/\Gamma(\chi_{c1} K^*(892)^0)$	$\Gamma_{245}/\Gamma_{249}$
VALUE (units $10^{-2}$ )	CL%

**20 ± 5 OUR FIT** Error includes scale factor of 1.1.  
**17.1±5.0±2.0** <sup>1</sup>AAIJ 13Ac LHCb  $pp$  at 7 TeV

<sup>1</sup> Uses  $B(\chi_{c1} \rightarrow J/\psi\gamma)/B(\chi_{c2} \rightarrow J/\psi\gamma) = 1.76 \pm 0.11$ .

$\Gamma(\chi_{c1} \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{246}/\Gamma$
VALUE (units $10^{-5}$ )	CL%

**1.12±0.25±0.12** <sup>1</sup>KUMAR 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1} K^0)/\Gamma_{\text{total}}$	$\Gamma_{247}/\Gamma$
VALUE (units $10^{-4}$ )	CL%

**3.93±0.27 OUR AVERAGE**  
 3.78 $^{+0.17}_{-0.16}$ ±0.33 <sup>1</sup>BHARDWAJ 11 BELL  $e^+e^- \rightarrow \Upsilon(4S)$   
 4.2 ± 0.3 ± 0.3 <sup>2</sup>AUBERT 09B BABR  $e^+e^- \rightarrow \Upsilon(4S)$   
 3.1  $^{+1.6}_{-1.1}$ ±0.1 <sup>3</sup>AVERY 00 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.51±0.33±0.45 <sup>1</sup>SONI 06 BELL Repl. by BHARDWAJ 11  
 4.53±0.41±0.51 <sup>1</sup>AUBERT 05J BABR Repl. by AUBERT 09B  
 4.3 ± 1.4 ± 0.2 <sup>4</sup>AUBERT 02 BABR Repl. by AUBERT 05J  
 <27 90 <sup>1</sup>ALAM 94 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses  $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .

<sup>3</sup> AVERY 00 reports  $(3.9^{+1.9}_{-1.3} \pm 0.4) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \chi_{c1} K^0)/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>4</sup> AUBERT 02 reports  $(5.4 \pm 1.4 \pm 1.1) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \chi_{c1} K^0)/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1} K^0)/\Gamma(J/\psi(1S) K^0)$	$\Gamma_{247}/\Gamma_{183}$
VALUE	CL%

**0.53±0.16±0.02** <sup>1</sup>AUBERT 02 BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 02 reports  $0.66 \pm 0.11 \pm 0.17$  from a measurement of  $[\Gamma(B^0 \rightarrow \chi_{c1} K^0)/\Gamma(B^0 \rightarrow J/\psi(1S) K^0)] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1} K^- \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{248}/\Gamma$
VALUE (units $10^{-4}$ )	CL%

**3.83±0.10±0.39** <sup>1</sup>MIZUK 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1} K^- \pi^+)/\Gamma(J/\psi(1S) K^- \pi^-)$	$\Gamma_{248}/\Gamma_{184}$
VALUE	CL%

**0.480±0.021±0.017** <sup>1</sup>LEES 12B BABR

<sup>1</sup> LEES 12B reports  $0.474 \pm 0.013 \pm 0.026$  from a measurement of  $[\Gamma(B^0 \rightarrow \chi_{c1} K^- \pi^+)/\Gamma(B^0 \rightarrow J/\psi(1S) K^- \pi^-)] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (34.4 \pm 1.5) \times 10^{-2}$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\chi_{c1} K^*(892)^0)/\Gamma_{\text{total}}$	$\Gamma_{249}/\Gamma$
VALUE (units $10^{-4}$ )	CL%

**2.39±0.19 OUR FIT** Error includes scale factor of 1.2.

**2.22 $^{+0.40}_{-0.31}$  OUR AVERAGE** Error includes scale factor of 1.6.

2.5 ± 0.2 ± 0.2 <sup>1</sup>AUBERT 09B BABR  $e^+e^- \rightarrow \Upsilon(4S)$   
 1.73 $^{+0.15}_{-0.12}$ ±0.34 <sup>2</sup>MIZUK 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.14±0.34±0.72 <sup>2</sup>SONI 06 BELL Repl. by MIZUK 08  
 3.27±0.42±0.64 <sup>2</sup>AUBERT 05J BABR Repl. by AUBERT 09B  
 3.9 ± 1.3 ± 0.1 <sup>3</sup>AUBERT 02 BABR Repl. by AUBERT 05J  
 <21 90 <sup>4</sup>ALAM 94 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $\chi_{c1,2} \rightarrow J/\psi\gamma$ . Assumes  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> AUBERT 02 reports  $(4.8 \pm 1.4 \pm 0.9) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \chi_{c1} K^*(892)^0)/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>4</sup> BORTOLETTO 92 assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\chi_{c1} K^*(892)^0)/\Gamma(J/\psi(1S) K^*(892)^0)$	$\Gamma_{249}/\Gamma_{185}$
VALUE (units $10^{-2}$ )	CL%

**18.7±1.5 OUR FIT** Error includes scale factor of 1.1.

**19.8±1.1±1.5** <sup>1</sup>AAIJ 13Ac LHCb  $pp$  at 7 TeV

<sup>1</sup> Uses  $B(\chi_{c1} \rightarrow J/\psi\gamma) = (34.4 \pm 1.5)\%$ .

$\Gamma(X(4051)^+ K^-, X^+ \rightarrow \chi_{c1} \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{250}/\Gamma$
VALUE (units $10^{-5}$ )	CL%

**3.0 $^{+1.5}_{-0.8}$ ±1.7** <sup>1</sup>MIZUK 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.8 90 <sup>1,2</sup>LEES 12B BABR

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses  $\chi_{c1} \rightarrow J/\psi\gamma$  mode. Uses  $\chi_{c1} \rightarrow J/\psi\gamma$  mode. Finds a good description of the data without this  $B^0 \rightarrow X(4051)^+ K^-$  decay mode in a fit.

$\Gamma(X(4248)^+ K^-, X^+ \rightarrow \chi_{c1} \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{251}/\Gamma$
VALUE (units $10^{-5}$ )	CL%

**4.0 $^{+2.3}_{-0.9}$ ±19.7** <sup>1</sup>MIZUK 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<4.0 90 <sup>1,2</sup>LEES 12B BABR

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses  $\chi_{c1} \rightarrow J/\psi\gamma$  mode. Finds a good description of the data without this  $B^0 \rightarrow X(4248)^+ K^-$  decay mode in a fit.

$\Gamma(\chi_{c1} K^*(892)^0)/\Gamma(\chi_{c1} K^0)$	$\Gamma_{249}/\Gamma_{247}$
VALUE	CL%

**0.72±0.11±0.12** AUBERT 05J BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.89±0.34±0.17 <sup>1</sup>AUBERT 02 BABR Repl. by AUBERT 05J

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{252}/\Gamma$
VALUE (units $10^{-6}$ )	CL%

**19.6 ± 0.5 OUR FIT**  
**19.6 ± 0.5 OUR AVERAGE**  
 20.00± 0.34±0.60 <sup>1</sup>DUH 13 BELL Repl. by DUH 13  
 19.1 ± 0.6 ± 0.6 <sup>1</sup>AUBERT 07B BABR  $e^+e^- \rightarrow \Upsilon(4S)$   
 18.0  $^{+2.3}_{-2.1}$ ±1.2 <sup>1</sup>BORNHEIM 03 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

19.9 ± 0.4 ± 0.8 <sup>1</sup>LIN 07A BELL Repl. by DUH 13  
 18.5 ± 1.0 ± 0.7 <sup>1</sup>CHAO 04 BELL Repl. by LIN 07A  
 17.9 ± 0.9 ± 0.7 <sup>1</sup>AUBERT 02Q BABR Repl. by AUBERT 07B  
 22.5 ± 1.9 ± 1.8 <sup>1</sup>CASEY 02 BELL Repl. by CHAO 04  
 19.3  $^{+3.4}_{-3.2}$ ±1.5 <sup>1</sup>ABE 01H BELL Repl. by CASEY 02  
 16.7 ± 1.6 ± 1.3 <sup>1</sup>AUBERT 01E BABR Repl. by AUBERT 02Q  
 < 66 90 <sup>2</sup>ABE 00C SLD  $e^+e^- \rightarrow Z$

<sup>1</sup> CRONIN-HEN...00 CLE2 Repl. by BORNHEIM 03

17.2  $^{+2.5}_{-2.4}$ ±1.2

15  $^{+5}_{-4}$ ±1.4

24  $^{+17}_{-11}$ ±2

< 17 90

< 30 90

< 90 90

< 81 90

< 26 90

<180 90

< 90 90

<320 90

<sup>3</sup> ADAM 96D DLPH  $e^+e^- \rightarrow Z$

<sup>4</sup> ASNER 96 CLE2 Sup. by ADAM 96D

<sup>5</sup> BUSKULIC 96V ALEP  $e^+e^- \rightarrow Z$

<sup>6</sup> ABREU 95N DLPH Sup. by ADAM 96D

<sup>7</sup> AKERS 94L OPAL  $e^+e^- \rightarrow Z$

<sup>8</sup> BATTLE 93 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>9</sup> ALBRECHT 91B ARG  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>10</sup> AVERY 89B CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>11</sup> AVERY 87 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

# Meson Particle Listings

$B^0$

- 1 Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .
- 2 ABE 00c assumes  $B(Z \rightarrow b\overline{b})=(21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0}=f_{B^+}=(39.7^{+1.8}_{-2.2}\%)$  and  $f_{B_s}=(10.5^{+1.8}_{-2.2}\%)$ .
- 3 ADAM 96d assumes  $f_{B^0}=f_{B^-}=0.39$  and  $f_{B_s}=0.12$ . Contributions from  $B^0$  and  $B_s$  decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral  $B$  mesons.
- 4 BUSKULIC 96v assumes PDG 96 production fractions for  $B^0$ ,  $B^+$ ,  $B_s$ ,  $b$  baryons.
- 5 Assumes a  $B^0$ ,  $B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12. Contributions from  $B^0$  and  $B_s^0$  decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral  $B$  mesons.
- 6 Assumes  $B(Z \rightarrow b\overline{b})=0.217$  and  $B_s^0$  ( $B_s^-$ ) fraction 39.5% (12%).
- 7 BATTLE 93 assumes equal production of  $B^0\overline{B}^0$  and  $B^+B^-$  at  $\Upsilon(4S)$ .
- 8 Assumes the  $\Upsilon(4S)$  decays 43% to  $B^0\overline{B}^0$ .

$\Gamma(K^+\pi^-)/\Gamma(K^0\pi^0)$		$\Gamma_{252}/\Gamma_{253}$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>2.16±0.16±0.16</b>	LIN	07A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.20 <sup>+0.50+0.22</sup> <sub>-0.58-0.32</sub>	<sup>1</sup> ABE	01H	BELL Repl. by LIN 07A
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$[\Gamma(K^+\pi^-)+\Gamma(\pi^+\pi^-)]/\Gamma_{total}$		$(\Gamma_{252}+\Gamma_{382})/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>19± 6 OUR AVERAGE</b>			
28 <sup>+15</sup> <sub>-10</sub> ±20	<sup>1</sup> ADAM	96D	DLPH $e^+e^- \rightarrow Z$
18 <sup>+6+3</sup> <sub>-5-4</sub>	17.2	ASNER	96 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
24 <sup>+8</sup> <sub>-7</sub> ±2	<sup>2</sup> BATTLE	93	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ADAM 96d assumes $f_{B^0}=f_{B^-}=0.39$ and $f_{B_s}=0.12$ . Contributions from $B^0$ and $B_s$ decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral $B$ mesons.			
<sup>2</sup> BATTLE 93 assumes equal production of $B^0\overline{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .			

$\Gamma(K^0\pi^0)/\Gamma_{total}$		$\Gamma_{253}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	CL%	DOCUMENT ID	TECN COMMENT
<b>9.9 ±0.5 OUR AVERAGE</b>			
9.68±0.46±0.50	<sup>1</sup> DUH	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
10.1 ±0.6 ±0.4	<sup>1</sup> LEES	13D	BABR $e^+e^- \rightarrow \Upsilon(4S)$
12.8 <sup>+4.0+1.7</sup> <sub>-3.3-1.4</sub>	<sup>1</sup> BORNHEIM	03	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
8.7 ±0.5 ±0.6	<sup>1</sup> FUJIKAWA	10A	BELL Repl. by DUH 13
10.3 ±0.7 ±0.6	<sup>1</sup> AUBERT	08E	BABR Repl. by LEES 13D
9.2 ±0.7 ±0.6	<sup>1</sup> LIN	07A	BELL Repl. by FUJIKAWA 10A
11.4 ±0.9 ±0.6	<sup>1</sup> AUBERT	05Y	BABR Repl. by AUBERT 08E
11.4 ±1.7 ±0.8	<sup>1</sup> AUBERT	04M	BABR Repl. by AUBERT 05Y
11.7 ±2.3 <sup>+1.2</sup> <sub>-1.3</sub>	<sup>1</sup> CHAO	04	BELL Repl. by LIN 07A
8.0 <sup>+3.3</sup> <sub>-3.1</sub> ±1.6	<sup>1</sup> CASEY	02	BELL Repl. by CHAO 04
16.0 <sup>+7.2+2.5</sup> <sub>-5.9-2.7</sub>	<sup>1</sup> ABE	01H	BELL Repl. by CASEY 02
8.2 <sup>+3.1</sup> <sub>-2.7</sub> ±1.2	<sup>1</sup> AUBERT	01E	BABR Repl. by AUBERT 04M
14.6 <sup>+5.9+2.4</sup> <sub>-5.1-3.3</sub>	<sup>1</sup> CRONIN-HEN..00	CLE2	Repl. by BORNHEIM 03
<41	90	GODANG	98 CLE2 Repl. by CRONIN-HENNESSY 00
<40	90	ASNER	96 CLE2 Rep. by GODANG 98
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta' K^0)/\Gamma_{total}$		$\Gamma_{254}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	DOCUMENT ID	TECN	COMMENT
<b>66 ± 4 OUR AVERAGE</b>	Error includes scale factor of 1.4.		
68.5± 2.2±3.1	<sup>1</sup> AUBERT	09AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$
58.9 <sup>+3.6</sup> <sub>-3.5</sub> ±4.3	<sup>1</sup> SCHUEMANN	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
89 <sup>+18</sup> <sub>-16</sub> ±9	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
66.6± 2.6±2.8	<sup>1</sup> AUBERT	07AE	BABR Repl. by AUBERT 09AV
67.4± 3.3±3.2	<sup>1</sup> AUBERT	05M	BABR AUBERT 07AE
60.6± 5.6±4.6	<sup>1</sup> AUBERT	03W	BABR Repl. by AUBERT 05M
55 <sup>+19</sup> <sub>-16</sub> ±8	<sup>1</sup> ABE	01M	BELL Repl. by SCHUEMANN 06
42 <sup>+13</sup> <sub>-11</sub> ±4	<sup>1</sup> AUBERT	01G	BABR Repl. by AUBERT 03W
47 <sup>+27</sup> <sub>-20</sub> ±9	BEHRENS	98	CLE2 Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta' K^*(892)^0)/\Gamma_{total}$		$\Gamma_{255}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	CL%	DOCUMENT ID	TECN COMMENT
<b>2.8±0.6 OUR AVERAGE</b>			
2.6±0.7±0.2	<sup>1</sup> SATO	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
3.1 <sup>+0.9</sup> <sub>-0.8</sub> ±0.3	<sup>1</sup> DEL-AMO-SA..10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.8±1.1±0.5	<sup>1</sup> AUBERT	07E	BABR Repl. by DEL-AMO-SANCHEZ 10A
< 2.6	90	<sup>1</sup> SCHUEMANN	07 BELL $e^+e^- \rightarrow \Upsilon(4S)$
< 7.6	90	<sup>1</sup> AUBERT,B	04D BABR Repl. by AUBERT 07E
<24	90	<sup>1</sup> RICHICHI	00 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<39	90	BEHRENS	98 CLE2 Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta' K_0^*(1430)^0)/\Gamma_{total}$		$\Gamma_{256}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	DOCUMENT ID	TECN	COMMENT
<b>6.3±1.3±0.9</b>	<sup>1</sup> DEL-AMO-SA..10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta' K_2^*(1430)^0)/\Gamma_{total}$		$\Gamma_{257}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	DOCUMENT ID	TECN	COMMENT
<b>13.7<sup>+3.0</sup><sub>-2.9</sub>±1.2</b>	<sup>1</sup> DEL-AMO-SA..10A	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta K^0)/\Gamma_{total}$		$\Gamma_{258}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	CL%	DOCUMENT ID	TECN COMMENT
<b>1.23±0.27<sub>-0.24</sub> OUR AVERAGE</b>			
1.27±0.33 <sub>-0.29</sub> ±0.08	<sup>1</sup> HOI	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
1.15 <sup>+0.43</sup> <sub>-0.38</sub> ±0.09	<sup>1</sup> AUBERT	09AV	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 1.9	90	<sup>1</sup> CHANG	07B BELL Repl. by HOI 12
< 2.9	90	<sup>1</sup> AUBERT,B	06V BABR $e^+e^- \rightarrow \Upsilon(4S)$
< 2.5	90	<sup>1</sup> AUBERT,B	05K BABR $e^+e^- \rightarrow \Upsilon(4S)$
< 2.0	90	<sup>1</sup> CHANG	05A BELL Repl. by CHANG 07B
< 5.2	90	<sup>1</sup> AUBERT	04H BABR Repl. by AUBERT,B 05K
< 9.3	90	<sup>1</sup> RICHICHI	00 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<33	90	BEHRENS	98 CLE2 Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta K^*(892)^0)/\Gamma_{total}$		$\Gamma_{259}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	CL%	DOCUMENT ID	TECN COMMENT
<b>15.9±1.0 OUR AVERAGE</b>			
15.2±1.2±1.0	<sup>1</sup> WANG	07B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
16.5±1.1±0.8	<sup>1</sup> AUBERT,B	06H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
13.6 <sup>+5.5</sup> <sub>-4.6</sub> ±1.6	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
18.6±2.3±1.2	<sup>1</sup> AUBERT,B	04D	BABR Repl. by AUBERT,B 06H
<30	90	BEHRENS	98 CLE2 Repl. by RICHICHI 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta K_0^*(1430)^0)/\Gamma_{total}$		$\Gamma_{260}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	DOCUMENT ID	TECN	COMMENT
<b>11.0±1.6±1.5</b>	<sup>1</sup> AUBERT,B	06H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\eta K_2^*(1430)^0)/\Gamma_{total}$		$\Gamma_{261}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	DOCUMENT ID	TECN	COMMENT
<b>9.6±1.8±1.1</b>	<sup>1</sup> AUBERT,B	06H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

$\Gamma(\omega K^0)/\Gamma_{total}$		$\Gamma_{262}/\Gamma$	
VALUE (units 10 <sup>-6</sup> )	CL%	DOCUMENT ID	TECN COMMENT
<b>4.8±0.4 OUR AVERAGE</b>			
4.5±0.4±0.3	<sup>1</sup> CHOBANOVA	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
5.4±0.8±0.3	<sup>1</sup> AUBERT	07AE	BABR $e^+e^- \rightarrow \Upsilon(4S)$
10.0 <sup>+5.4</sup> <sub>-4.2</sub> ±1.4	<sup>1</sup> JESSOP	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6.2±1.0±0.4	<sup>1</sup> AUBERT,B	06E	BABR Repl. by AUBERT 07AE
4.4 <sup>+0.8</sup> <sub>-0.7</sub> ±0.4	<sup>1</sup> JEN	06	BELL Repl. by CHOBANOVA 14
5.9 <sup>+1.6</sup> <sub>-1.3</sub> ±0.5	<sup>1</sup> AUBERT	04H	BABR Repl. by AUBERT,B 06E
4.0 <sup>+1.9</sup> <sub>-1.6</sub> ±0.5	<sup>1</sup> WANG	04A	BELL Repl. by JEN 06
<13	90	<sup>1</sup> AUBERT	01G BABR Repl. by AUBERT 04H
<57	90	<sup>1</sup> BERGFELD	98 CLE2 Repl. by JESSOP 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

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Meson Particle Listings  
 $B^0$  $\Gamma(a_0(980)^0 K^0, a_0^0 \rightarrow \eta \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{263}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.8</b>	90	<sup>1</sup> AUBERT,BE	04 BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of charged and neutral $B$ mesons at $\gamma(4S)$ .				

 $\Gamma(b_1^0 K^0, b_1^0 \rightarrow \omega \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{264}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.8</b>	90	<sup>1</sup> AUBERT	08AQ BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(a_0(980)^\pm K^\mp, a_0^\pm \rightarrow \eta \pi^\pm)/\Gamma_{\text{total}}$   $\Gamma_{265}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.9</b>	90	<sup>1</sup> AUBERT	07Y BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.1	90	<sup>1</sup> AUBERT,BE	04 BABR	Repl. by AUBERT 07Y
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(b_1^- K^+, b_1^- \rightarrow \omega \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{266}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>7.4 \pm 1.0 \pm 1.0</math></b>		<sup>1</sup> AUBERT	07BI BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(b_1^0 K^{*0}, b_1^0 \rightarrow \omega \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{267}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;8.0 \times 10^{-6}</math></b>	90	<sup>1</sup> AUBERT	09AF BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(b_1^- K^{*+}, b_1^- \rightarrow \omega \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{268}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5.0 \times 10^{-6}</math></b>	90	<sup>1</sup> AUBERT	09AF BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(a_0(1450)^\pm K^\mp, a_0^\pm \rightarrow \eta \pi^\pm)/\Gamma_{\text{total}}$   $\Gamma_{269}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.1</b>	90	<sup>1</sup> AUBERT	07Y BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(K_S^0 X^0 (\text{Familon}))/\Gamma_{\text{total}}$   $\Gamma_{270}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;53</b>	90	<sup>1</sup> AMMAR	01B CLE2	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> AMMAR 01B searched for the two-body decay of the $B$ meson to a massless neutral feebly-interacting particle $X^0$ such as the familon, the Nambu-Goldstone boson associated with a spontaneously broken global family symmetry.				

 $\Gamma(\omega K^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{271}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.0 \pm 0.5</math> OUR AVERAGE</b>				
$2.2 \pm 0.6 \pm 0.2$		<sup>1</sup> AUBERT	09H BABR	$e^+ e^- \rightarrow \gamma(4S)$
$1.8 \pm 0.7 \pm 0.3$		<sup>1</sup> GOLDENZWE..08	BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 4.2	90	<sup>1</sup> AUBERT,B	06T BABR	Repl. by AUBERT 09H
< 6.0	90	<sup>1</sup> AUBERT	05o BABR	Repl. by AUBERT,B 06T
<23	90	<sup>1</sup> BERGFELD	98 CLE2	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(\omega (K\pi)_0^{*0})/\Gamma_{\text{total}}$   $\Gamma_{272}/\Gamma$ 

$(K\pi)_0^{*0}$ is the total S-wave composed of $K_0^*(1430)$ and nonresonant that are described using LASS shape.				
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>18.4 \pm 1.8 \pm 1.7</math></b>		<sup>1</sup> AUBERT	09H BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(\omega K_2^*(1430)^0)/\Gamma_{\text{total}}$   $\Gamma_{273}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>16.0 \pm 1.6 \pm 3.0</math></b>		<sup>1</sup> AUBERT	09H BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(\omega K_2^*(1430)^0)/\Gamma_{\text{total}}$   $\Gamma_{274}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>10.1 \pm 2.0 \pm 1.1</math></b>		<sup>1</sup> AUBERT	09H BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

 $\Gamma(\omega K^+ \pi^- \text{nonresonant})/\Gamma_{\text{total}}$   $\Gamma_{275}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.1 \pm 0.7 \pm 0.7</math></b>		<sup>1,2</sup> GOLDENZWE..08	BELL	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> For the $K\pi$ mass range 0.755–1.250 GeV/c <sup>2</sup> , excluding $K^*(892)$ .				

 $\Gamma(K^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{276}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>37.8 \pm 3.2</math> OUR AVERAGE</b>				
$38.5 \pm 1.0 \pm 3.9$		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
$36.6^{+4.2}_{-4.3} \pm 3.0$		<sup>1</sup> CHANG	04 BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$35.7^{+2.6}_{-1.5} \pm 2.2$		<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<40	90	<sup>1</sup> ECKHART	02 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.				

 $\Gamma(K^+ \rho^-)/\Gamma_{\text{total}}$   $\Gamma_{277}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>7.0 \pm 0.9</math> OUR AVERAGE</b>				
$6.6 \pm 0.5 \pm 0.8$		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
$15.1^{+3.4+2.4}_{-3.3-2.6}$		<sup>1</sup> CHANG	04 BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.0^{+0.8}_{-1.3} \pm 0.6$		<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
$7.3^{+1.3}_{-1.2} \pm 1.3$		<sup>1</sup> AUBERT	03T BABR	Repl. by AUBERT 08AQ
<32	90	<sup>1</sup> JESSOP	00 CLE2	$e^+ e^- \rightarrow \gamma(4S)$
<35	90	ASNER	96 CLE2	Repl. by JESSOP 00
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.				

 $\Gamma(K^+ \rho(1450)^-)/\Gamma_{\text{total}}$   $\Gamma_{278}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.4 \pm 1.0 \pm 0.6</math></b>		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2.1	90	<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.				

 $\Gamma(K^+ \rho(1700)^-)/\Gamma_{\text{total}}$   $\Gamma_{279}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.6 \pm 0.6 \pm 0.4</math></b>		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.1	90	<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.				

 $\Gamma((K^+ \pi^- \pi^0) \text{non-resonant})/\Gamma_{\text{total}}$   $\Gamma_{280}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>2.8 \pm 0.5 \pm 0.4</math></b>		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.4 \pm 0.9 \pm 0.5$		<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<9.4	90	<sup>1</sup> CHANG	04 BELL	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays. The quoted value is only for the flat part of the non-resonant component.				

 $\Gamma((K\pi)_0^{*+} \pi^-, (K\pi)_0^{*+} \rightarrow K^+ \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{281}/\Gamma$ 

$(K\pi)_0^{*+}$  is the total S-wave composed of  $K_0^*(1430)$  and nonresonant that are described using LASS shape.

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>34.2 \pm 2.4 \pm 4.1</math></b>		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$9.4^{+1.1+2.3}_{-1.3-2.1}$		<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.				

 $\Gamma((K\pi)_0^{*0} \pi^0, (K\pi)_0^{*0} \rightarrow K^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{282}/\Gamma$ 

$(K\pi)_0^{*0}$  is the total S-wave composed of  $K_0^*(1430)$  and nonresonant that are described using LASS shape.

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>8.6 \pm 1.1 \pm 1.3</math></b>		<sup>1,2</sup> LEES	11 BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.7^{+1.1+2.8}_{-0.9-2.6}$		<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.				

 $\Gamma(K_2^*(1430)^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{283}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.0</b>	90	<sup>1</sup> AUBERT	08AQ BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				

# Meson Particle Listings

$B^0$

$\Gamma(K^*(1680)^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{284}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;7.5</b>	90	<sup>1</sup> AUBERT	08AQ	BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K_X^{*0}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{285}/\Gamma$
$K_X^{*0}$ stands for the possible candidates of $K^*(1410)$ , $K_0^*(1430)$ and $K_2^*(1430)$ .					
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>6.1 \pm 1.6 + 0.5</math> <math>-1.5 - 0.6</math></b>		<sup>1</sup> CHANG	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^0\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{286}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>52.0 \pm 2.4</math> OUR FIT</b>	Error includes scale factor of 1.3.				
<b><math>49.6 \pm 2.0</math> OUR AVERAGE</b>					
50.2 $\pm$ 1.5 $\pm$ 1.8		<sup>1</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
47.5 $\pm$ 2.4 $\pm$ 3.7		<sup>2</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
50 $-10$ $-9$ $\pm$ 7		<sup>1</sup> ECKHART	02	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
43.0 $\pm$ 2.3 $\pm$ 2.3		<sup>1</sup> AUBERT	06i	BABR Repl. by AUBERT 09AU	
43.7 $\pm$ 3.8 $\pm$ 3.4		<sup>1</sup> AUBERT,B	04o	BABR Repl. by AUBERT 06i	
45.4 $\pm$ 5.2 $\pm$ 5.9		<sup>1</sup> GARMASH	04	BELL Repl. by GARMASH 07	
<440	90	ALBRECHT	91E	ARG $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K^0\pi^+\pi^-$  final state decays.

$\Gamma(K^0\pi^+\pi^- \text{ non-resonant})/\Gamma_{\text{total}}$					$\Gamma_{287}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>14.7 \pm 4.0</math> OUR AVERAGE</b>	Error includes scale factor of 2.1.				
$11.1 \pm 2.5 \pm 0.9$		<sup>1</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
19.9 $\pm$ 2.5 $+1.7$ $-2.0$		<sup>2</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Uses Dalitz plot analysis of the $B^0 \rightarrow K^0\pi^+\pi^-$ final state decays.					

$\Gamma(K^0\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{288}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>4.7 \pm 0.6</math> OUR AVERAGE</b>					
4.4 $\pm$ 0.7 $\pm$ 0.3		<sup>1</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
6.1 $\pm$ 1.0 $+1.1$ $-1.2$		<sup>2</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
4.9 $\pm$ 0.8 $\pm$ 0.9		<sup>1</sup> AUBERT	07f	BABR Repl. by AUBERT 09AU	
< 39	90	ASNER	96	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	
< 320	90	ALBRECHT	91B	ARG $e^+e^- \rightarrow \Upsilon(4S)$	
< 500	90	<sup>3</sup> AVERY	89B	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	
<64000	90	<sup>4</sup> AVERY	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K^0\pi^+\pi^-$  final state decays.

<sup>3</sup> AVERY 89B reports  $< 5.8 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

<sup>4</sup> AVERY 87 reports  $< 0.08$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K^*(892)^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{289}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>8.4 \pm 0.8</math> OUR AVERAGE</b>					
8.0 $\pm$ 1.1 $\pm$ 0.8		<sup>1,2</sup> LEES	11	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
8.3 $\pm$ 0.9 $\pm$ 0.8		<sup>2,3</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
8.4 $\pm$ 1.1 $+1.0$ $-0.9$		<sup>3</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
16 $+6$ $-5$ $\pm$ 2		<sup>2</sup> ECKHART	02	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
12.6 $\pm$ 2.7 $\pm$ 0.9		<sup>1,2</sup> AUBERT	08AQ	BABR Repl. by LEES 11	
11.0 $\pm$ 1.5 $\pm$ 0.71		<sup>2</sup> AUBERT	06i	BABR Repl. by AUBERT 09AU	
12.9 $\pm$ 2.4 $\pm$ 1.4		<sup>2</sup> AUBERT,B	04o	BABR Repl. by AUBERT 06i	
14.8 $\pm$ 4.6 $+2.8$ $-4.4 - 1.3$		<sup>2</sup> CHANG	04	BELL Repl. by GARMASH 07	
< 72	90	ASNER	96	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
<620	90	ALBRECHT	91B	ARG $e^+e^- \rightarrow \Upsilon(4S)$	
<380	90	<sup>4</sup> AVERY	89B	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	
<560	90	<sup>5</sup> AVERY	87	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses Dalitz plot analysis of  $B^0 \rightarrow K^+\pi^-\pi^0$  decays.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K^0\pi^+\pi^-$  final state decays.

<sup>4</sup> AVERY 89B reports  $< 4.4 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

<sup>5</sup> AVERY 87 reports  $< 7 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K_0^*(1430)^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{290}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>33 \pm 7</math> OUR AVERAGE</b>	Error includes scale factor of 2.0.				
29.9 $-2.3$ $-1.7$ $\pm$ 3.6		<sup>1,2</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>2</sup> GARMASH 07 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

49.7  $\pm$  3.8  $+6.8$   
 $-8.2$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K^0\pi^+\pi^-$  final state decays.

$\Gamma(K_X^{*+}\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{291}/\Gamma$
$K_X^{*+}$ stands for the possible candidates of $K^*(1410)$ , $K_0^*(1430)$ and $K_2^*(1430)$ .					
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>5.1 \pm 1.5 + 0.6</math> <math>-0.7</math></b>		<sup>1</sup> CHANG	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(1410)^+\pi^-, K^{*+} \rightarrow K^0\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{292}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.8</b>	90	<sup>1</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K^0\pi^+\pi^-$  final state decays.

$\Gamma(f_0(980)K^0, f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{293}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>7.0 \pm 0.9</math> OUR AVERAGE</b>					
6.9 $\pm$ 0.8 $\pm$ 0.6		<sup>1</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
7.6 $\pm$ 1.7 $+0.9$ $-1.3$		<sup>2</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
5.5 $\pm$ 0.7 $\pm$ 0.6		<sup>1</sup> AUBERT	06i	BABR Repl. by AUBERT 09AU	
<360	90	<sup>3</sup> AVERY	89B	CLEO $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K^0\pi^+\pi^-$  final state decays.

<sup>3</sup> AVERY 89B reports  $< 4.2 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(f_2(1270)K^0)/\Gamma_{\text{total}}$					$\Gamma_{294}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.7 \pm 1.0 + 0.8</math> <math>-0.9</math></b>		<sup>1</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
<2.5	90	<sup>2</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> GARMASH 07 reports  $B(B^0 \rightarrow f_2(1270)K^0) \times B(f_2(1270) \rightarrow \pi^+\pi^-) < 1.4 \times 10^{-6}$  using Dalitz plot analysis. We compute  $B(B^0 \rightarrow f_2(1270)K^0)$  using the PDG value  $B(f_2(1270) \rightarrow \pi\pi) = 84.2 \times 10^{-2}$  and 2/3 for the  $\pi^+\pi^-$  fraction.

$\Gamma(f_X(1300)K^0, f_X \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{295}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>1.81 \pm 0.55 \pm 0.48</math></b>		<sup>1</sup> AUBERT	09AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{296}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.3 \pm 0.5 \pm 0.4</math></b>		<sup>1,2</sup> LEES	11	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
3.6 $\pm$ 0.7 $\pm$ 0.4		<sup>1,2</sup> AUBERT	08AQ	BABR Repl. by LEES 11	
< 3.5	90	<sup>2</sup> CHANG	04	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
< 3.6	90	JESSOP	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
<28	90	ASNER	96	CLE2 Repl. by JESSOP 00	

<sup>1</sup> Uses Dalitz plot analysis of  $B^0 \rightarrow K^+\pi^-\pi^0$  decays.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K_2^*(1430)^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{297}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 6</b>	90	<sup>1</sup> GARMASH	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$	
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
< 16.2	90	<sup>2,3</sup> AUBERT	08AQ	BABR $e^+e^- \rightarrow \Upsilon(4S)$	
< 18	90	<sup>3</sup> GARMASH	04	BELL Repl. by GARMASH 07	
<2600	90	ALBRECHT	91B	ARG $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> GARMASH 07 reports  $B(B^0 \rightarrow K_2^*(1430)^+\pi^-) \times B(K_2^{*+} \rightarrow K^0\pi^+) < 2.1 \times 10^{-6}$  using Dalitz plot analysis. We compute  $B(B^0 \rightarrow K_2^*(1430)^+\pi^-)$  using the PDG value  $B(K_2^*(1430) \rightarrow K\pi) = 49.9 \times 10^{-2}$  and 2/3 for the  $K^0\pi^+$  fraction.

<sup>2</sup> Uses Dalitz plot analysis of  $B^0 \rightarrow K^+\pi^-\pi^0$  decays.

<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(1680)^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{298}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;10</b>	90	<sup>1</sup> GARMASH 07	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
	90	<sup>2,3</sup> AUBERT 08AQ	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> GARMASH 07 reports $B(B^0 \rightarrow K^*(1680)^+\pi^-) \times B(K^* \rightarrow K^0\pi^+) < 2.6 \times 10^{-6}$ using Dalitz plot analysis. We compute $B(B^0 \rightarrow K^*(1680)^+\pi^-)$ using the PDG value $B(K^*(1680) \rightarrow K\pi) = 38.7 \times 10^{-2}$ and 2/3 for the $K^0\pi^+$ fraction.					
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+\pi^-\pi^0$ decays.					
<sup>3</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{299}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.3 <math>\times 10^{-4}</math></b>	90	<sup>1</sup> ADAM 96D	DLPH	$e^+e^- \rightarrow Z$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>&lt;2.1 <math>\times 10^{-4}</math></b>	90	<sup>2</sup> ABREU 95N	DLPH	Sup. by ADAM 96D	
<sup>1</sup> ADAM 96D assumes $f_{B^0} = f_{B^-} = 0.39$ and $f_{B_s} = 0.12$ . Contributions from $B^0$ and $B_s$ decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral $B$ mesons.					
<sup>2</sup> Assumes a $B^0$ , $B^-$ production fraction of 0.39 and a $B_s$ production fraction of 0.12. Contributions from $B^0$ and $B_s$ decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral $B$ mesons.					

$\Gamma(\rho^0 K^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{300}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>2.8 <math>\pm 0.5 \pm 0.5</math></b>		<sup>1,2</sup> KYEONG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Required $0.75 < m_{K^+\pi^-} < 1.20$ GeV/ $c^2$ .					

$\Gamma(f_0(980)K^+\pi^-, f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$					$\Gamma_{301}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>1.4 <math>\pm 0.4 \pm 0.3</math> -0.4</b>		<sup>1,2</sup> KYEONG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Required $0.75 < m_{K^+\pi^-} < 1.2$ GeV/ $c^2$ .					

$\Gamma(K^+\pi^-\pi^+\pi^- \text{ nonresonant})/\Gamma_{\text{total}}$					$\Gamma_{302}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.1 <math>\times 10^{-6}</math></b>	90	<sup>1,2</sup> KYEONG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Required $0.55 < m_{\pi^+\pi^-} < 1.42$ and $0.75 < m_{K^+\pi^-} < 1.20$ GeV/ $c^2$ .					

$\Gamma(K^*(892)^0\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{303}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>54.5 <math>\pm 2.9 \pm 4.3</math></b>		<sup>1</sup> AUBERT 07AS	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>4.5 <math>\pm 1.1 \pm 0.9</math> -1.0 -1.6</b>		<sup>1,2</sup> KYEONG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>&lt;1400</b>	90	ALBRECHT 91E	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Required $0.55 < m_{\pi^+\pi^-} < 1.42$ GeV/ $c^2$ .					

$\Gamma(K^*(892)^0\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{304}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>3.9 <math>\pm 1.3</math> OUR AVERAGE</b>		Error includes scale factor of 1.9.			
<b>5.1 <math>\pm 0.6 \pm 0.6</math> -0.8</b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>2.1 <math>\pm 0.8 \pm 0.9</math> -0.7 -0.5</b>		<sup>1</sup> KYEONG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>5.6 <math>\pm 0.9 \pm 1.3</math></b>		<sup>1</sup> AUBERT,B 06G	BABR	Repl. by LEES 12K	
<b>&lt; 34</b>	90	<sup>2</sup> GODANG 02	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>&lt;286</b>	90	<sup>3</sup> ABE 00C	SLD	$e^+e^- \rightarrow Z$	
<b>&lt;460</b>	90	ALBRECHT 91B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>&lt;580</b>	90	<sup>4</sup> AVERY 89B	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>&lt;960</b>	90	<sup>5</sup> AVERY 87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to $2.4 \times 10^{-5}$ .					
<sup>3</sup> ABE 00C assumes $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$ and the $B$ fractions $f_{B^0} = f_{B^+} = (39.7 \pm 1.8)\%$ and $f_{B_s} = (10.5 \pm 1.8)\%$ .					
<sup>4</sup> AVERY 89B reports $< 6.7 \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.					
<sup>5</sup> AVERY 87 reports $< 1.2 \times 10^{-3}$ assuming the $\Upsilon(4S)$ decays 40% to $B^0\bar{B}^0$ . We rescale to 50%.					

$\Gamma(K^*(892)^0 f_0(980), f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$					$\Gamma_{305}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>3.9 <math>\pm 2.1</math> -1.8</b> OUR AVERAGE		Error includes scale factor of 3.9.			
<b>5.7 <math>\pm 0.6 \pm 0.4</math></b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>1.4 <math>\pm 0.6 \pm 0.6</math> -0.5 -0.4</b>		<sup>1,2</sup> KYEONG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>&lt; 4.3</b>	90	<sup>1</sup> AUBERT,B 06G	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<b>&lt;170</b>	90	<sup>3</sup> AVERY 89B	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> The upper limit is $2.2 \times 10^{-6}$ at 90% CL.					
<sup>3</sup> AVERY 89B reports $< 2.0 \times 10^{-4}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.					

$\Gamma(K_1(1270)^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{306}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.0 <math>\times 10^{-5}</math></b>	90	<sup>1</sup> AUBERT 10D	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_1(1400)^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{307}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.7 <math>\times 10^{-5}</math></b>	90	<sup>1</sup> AUBERT 10D	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>&lt;1.1 <math>\times 10^{-3}</math></b>	90	ALBRECHT 91B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(a_1(1260)^-K^+)/\Gamma_{\text{total}}$					$\Gamma_{308}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>16.3 <math>\pm 2.9 \pm 2.3</math></b>		<sup>1,2</sup> AUBERT 08F	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>&lt;230</b>	90	<sup>3</sup> ADAM 96D	DLPH	$e^+e^- \rightarrow Z$	
<b>&lt;390</b>	90	<sup>4</sup> ABREU 95N	DLPH	Sup. by ADAM 96D	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Assumes $a_1^\pm$ decays only to $3\pi$ and $B(a_1^\pm \rightarrow \pi^\pm\pi^\mp\pi^\pm) = 0.5$ .					
<sup>3</sup> ADAM 96D assumes $f_{B^0} = f_{B^-} = 0.39$ and $f_{B_s} = 0.12$ . Contributions from $B^0$ and $B_s$ decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral $B$ mesons.					
<sup>4</sup> Assumes a $B^0$ , $B^-$ production fraction of 0.39 and a $B_s$ production fraction of 0.12. Contributions from $B^0$ and $B_s$ decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral $B$ mesons.					

$\Gamma(K^*(892)^+\rho^-)/\Gamma_{\text{total}}$					$\Gamma_{309}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>10.3 <math>\pm 2.3 \pm 1.3</math></b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<b>&lt;12.0</b>	90	<sup>1</sup> AUBERT,B 06G	BABR	Repl. by LEES 12K	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^+\rho^-)/\Gamma_{\text{total}}$					$\Gamma_{310}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>28 <math>\pm 10 \pm 6</math></b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_1(1400)^0\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{311}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.0 <math>\times 10^{-3}</math></b>	90	ALBRECHT 91B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(K_0^*(1430)^0\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{312}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>27 <math>\pm 4 \pm 4</math></b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^0 f_0(980), f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$					$\Gamma_{313}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>2.7 <math>\pm 0.7 \pm 0.6</math></b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_2^*(1430)^0 f_0(980), f_0 \rightarrow \pi\pi)/\Gamma_{\text{total}}$					$\Gamma_{314}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>8.6 <math>\pm 1.7 \pm 1.0</math></b>		<sup>1</sup> LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

# Meson Particle Listings

$B^0$

## $\Gamma(K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_{315}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.13±0.05 OUR AVERAGE</b>				
0.10±0.08±0.04		<sup>1,2</sup> DUH 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.12 $^{+0.08}_{-0.07}$ ±0.01		<sup>3</sup> AAIJ 12AR	LHCB	$pp$ at 7 TeV
0.23±0.10±0.10		<sup>4</sup> AALTONEN 12L	CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.7	90	<sup>5</sup> AALTONEN 09c	CDF	Repl. by AALTONEN 12L
< 0.5	90	<sup>2</sup> AUBERT 07B	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 0.41	90	<sup>2</sup> LIN 07	BELL	Repl. by DUH 13
< 1.8	90	<sup>6</sup> ABULENCIA,A 06D	CDF	Repl. by AALTONEN 09c
< 0.37	90	ABE 05G	BELL	Repl. by LIN 07
< 0.7	90	CHAO 04	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 0.8	90	<sup>2</sup> BORNHEIM 03	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 0.6	90	<sup>2</sup> AUBERT 02Q	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 0.9	90	<sup>2</sup> CASEY 02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 2.7	90	<sup>2</sup> ABE 01H	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 2.5	90	<sup>2</sup> AUBERT 01E	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 66	90	<sup>7</sup> ABE 00c	SLD	$e^+e^- \rightarrow Z$
< 1.9	90	<sup>2</sup> CRONIN-HEN..00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 4.3	90	GODANG 98	CLE2	Repl. by CRONIN-HENNESSY 00
< 46		<sup>8</sup> ADAM 96D	DLPH	$e^+e^- \rightarrow Z$
< 4	90	ASNER 96	CLE2	Repl. by GODANG 98
< 18	90	<sup>9</sup> BUSKULIC 96v	ALEP	$e^+e^- \rightarrow Z$
<120	90	<sup>10</sup> ABREU 95N	DLPH	Sup. by ADAM 96D
< 7	90	<sup>2</sup> BATTLE 93	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> DUH 13 reports also for the same data  $B(B^0 \rightarrow K^+K^-) < 0.20 \times 10^{-6}$  at 90% CL.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> AAIJ 12AR reports  $[\Gamma(B^0 \rightarrow K^+K^-)/\Gamma_{\text{total}}] / [B(B^0_s \rightarrow K^+K^-)] / [\Gamma(\bar{B} \rightarrow B^0_s)/\Gamma(\bar{B} \rightarrow B^0)] = 0.018^{+0.008}_{-0.007} \pm 0.009$  which we multiply by our best values  $B(B^0_s \rightarrow K^+K^-) = (2.52 \pm 0.17) \times 10^{-5}$ ,  $\Gamma(\bar{B} \rightarrow B^0_s)/\Gamma(\bar{B} \rightarrow B^0) = 0.256 \pm 0.014$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.  
<sup>4</sup> Reported a central value of  $(0.23 \pm 0.10 \pm 0.10) \times 10^{-6}$  using  $B(B^0 \rightarrow K^+\pi^-) = (19.4 \pm 0.6) \times 10^{-6}$ .  
<sup>5</sup> Obtains this result from  $B(K^+K^-)/B(K^+\pi^-) = 0.020 \pm 0.008 \pm 0.006$ , assuming  $B(B^0 \rightarrow K^+\pi^-) = (19.4 \pm 0.6) \times 10^{-6}$ .  
<sup>6</sup> ABULENCIA,A 06D obtains this from  $\Gamma(K^+K^-)/\Gamma(K^+\pi^-) < 0.10$  at 90% CL, assuming  $B(B^0 \rightarrow K^+\pi^-) = (18.9 \pm 0.7) \times 10^{-6}$ .  
<sup>7</sup> ABE 00c assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$  and  $f_{B_s} = (10.5^{+1.8}_{-2.2})\%$ .  
<sup>8</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ . Contributions from  $B^0$  and  $B_s$  decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral  $B$  mesons.  
<sup>9</sup> BUSKULIC 96v assumes PDG 96 production fractions for  $B^0$ ,  $B^+$ ,  $B_s$ ,  $b$  baryons.  
<sup>10</sup> Assumes a  $B^0$ ,  $B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12. Contributions from  $B^0$  and  $B_s$  decays cannot be separated. Limits are given for the weighted average of the decay rates for the two neutral  $B$  mesons.

## $\Gamma(K^0\bar{K}^0)/\Gamma_{\text{total}}$ $\Gamma_{316}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.21±0.16 OUR AVERAGE</b>				
1.26±0.19±0.05		<sup>1</sup> DUH 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
1.08±0.28±0.11		<sup>1</sup> AUBERT,BE 06c	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.87 $^{+0.25}_{-0.20}$ ±0.09		<sup>1</sup> LIN 07	BELL	Repl. by DUH 13
0.8 ± 0.3 ± 0.9		<sup>1</sup> ABE 05G	BELL	Repl. by LIN 07
1.19 $^{+0.40}_{-0.35}$ ±0.13		<sup>1</sup> AUBERT,BE 05E	BABR	Repl. by AUBERT,BE 06c
< 1.8	90	<sup>1</sup> AUBERT 04M	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 1.5	90	<sup>1</sup> CHAO 04	BELL	Repl. by ABE 05G
< 3.3	90	<sup>1</sup> BORNHEIM 03	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 4.1	90	<sup>1</sup> CASEY 02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<17	90	GODANG 98	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K^0K^-\pi^+)/\Gamma_{\text{total}}$ $\Gamma_{317}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>6.5±0.8 OUR FIT</b>				
<b>6.4±1.0±0.6</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<18	90	<sup>1</sup> GARMASH 04	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<21	90	<sup>1</sup> ECKHART 02	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K^*(892)^\pm K^\mp)/\Gamma_{\text{total}}$ $\Gamma_{318}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.4 × 10<sup>-6</sup></b>	90	AAIJ 14BM	LHCB	$pp$ at 7 TeV

## $\Gamma(K^0K^-\pi^+)/\Gamma(K^0\pi^+\pi^-)$ $\Gamma_{317}/\Gamma_{286}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.126±0.015 OUR FIT</b>			
<b>0.128±0.017±0.009</b>	AAIJ	13BP LHCB	$pp$ at 7 TeV

## $[\Gamma(\bar{K}^{*0}K^0) + \Gamma(K^{*0}\bar{K}^0)]/\Gamma_{\text{total}}$ $\Gamma_{319}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.96</b>	90	<sup>1</sup> AAIJ 16	LHCB	$pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.9	90	<sup>2</sup> AUBERT,BE 06N	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes  $B(B^0 \rightarrow K^0\pi^+\pi^-) = (4.96 \pm 0.20) \times 10^{-5}$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K^+K^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{320}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.17±0.60±0.24</b>		<sup>1</sup> GAUR 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<19	90	<sup>1</sup> ECKHART 02	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K_S^0K_S^0\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{321}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.9 × 10<sup>-6</sup></b>	90	<sup>1</sup> AUBERT 09AD	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K_S^0K_S^0\eta)/\Gamma_{\text{total}}$ $\Gamma_{322}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0 × 10<sup>-6</sup></b>	90	<sup>1</sup> AUBERT 09AD	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K_S^0K_S^0\eta')/\Gamma_{\text{total}}$ $\Gamma_{323}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.0 × 10<sup>-6</sup></b>	90	<sup>1</sup> AUBERT 09AD	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(K^0K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_{324}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>24.9±3.1 OUR FIT</b>				Error includes scale factor of 3.0.
<b>26.6±1.2 OUR AVERAGE</b>				
26.5±0.9±0.8		<sup>1,2</sup> LEES 12o	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
28.3±3.3±4.0		<sup>1</sup> GARMASH 04	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
23.8±2.0±1.6		<sup>1</sup> AUBERT,B 04v	BABR	Repl. by LEES 12o
<1300	90	ALBRECHT 91E	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> All intermediate charmonium and charm resonances are removed, except of  $\chi_{c0}$ .

## $\Gamma(K^0K^+K^-)/\Gamma(K^0\pi^+\pi^-)$ $\Gamma_{324}/\Gamma_{286}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.48 ± 0.06 OUR FIT</b>			Error includes scale factor of 2.7.
<b>0.385 ± 0.031 ± 0.023</b>	AAIJ	13BP LHCB	$pp$ at 7 TeV

## $\Gamma(K^0\phi)/\Gamma_{\text{total}}$ $\Gamma_{325}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>7.3±0.7 OUR AVERAGE</b>				
7.1 $^{+0.6}_{-0.3}$ ±0.4		<sup>1</sup> LEES 12o	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
9.0 $^{+2.2}_{-1.8}$ ±0.7		<sup>1</sup> CHEN 03B	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8.4 $^{+1.5}_{-1.3}$ ±0.5		<sup>1</sup> AUBERT 04A	BABR	Repl. by LEES 12o
8.1 $^{+3.1}_{-2.5}$ ±0.8		<sup>1</sup> AUBERT 01D	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 12.3	90	<sup>1</sup> BRIERE 01	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 31	90	<sup>1</sup> BERGFELD 98	CLE2	
< 88	90	ASNER 96	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 720	90	ALBRECHT 91B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
< 420	90	<sup>2</sup> AVERY 89B	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<1000	90	<sup>3</sup> AVERY 87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AVERY 89B reports  $< 4.9 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.  
<sup>3</sup> AVERY 87 reports  $< 1.3 \times 10^{-3}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0\bar{B}^0$ . We rescale to 50%.

## $\Gamma(f_0(980)K^0, f_0 \rightarrow K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_{326}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>7.0<math>^{+2.6}_{-1.8}</math>±2.4</b>	<sup>1</sup> LEES 12o	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

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Meson Particle Listings  
 $B^0$ 

$\Gamma(f_0(1500)K^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{327}/\Gamma$
VALUE (units $10^{-6}$ )				

**$13.3^{+5.8}_{-4.4} \pm 3.2$**  <sup>1</sup> LEES 12o BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(f'_2(1525)^0K^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{328}/\Gamma$
VALUE (units $10^{-6}$ )				

**$0.29^{+0.27}_{-0.18} \pm 0.36$**  <sup>1</sup> LEES 12o BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(f_0(1710)K^0, f_0 \rightarrow K^+K^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{329}/\Gamma$
VALUE (units $10^{-6}$ )				

**$4.4 \pm 0.7 \pm 0.5$**  <sup>1</sup> LEES 12o BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^0K^+K^- \text{ nonresonant})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{330}/\Gamma$
VALUE (units $10^{-6}$ )				

**$33 \pm 5 \pm 9$**  <sup>1</sup> LEES 12o BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K_S^0K_S^0K_S^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{331}/\Gamma$
VALUE (units $10^{-6}$ )				

**$6.0 \pm 0.5$  OUR AVERAGE** Error includes scale factor of 1.1.

$6.19 \pm 0.48 \pm 0.19$  <sup>1</sup> LEES 12i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

$4.2^{+1.6}_{-1.3} \pm 0.8$  <sup>1</sup> GARMASH 04 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$6.9^{+0.9}_{-0.8} \pm 0.6$  <sup>1</sup> AUBERT,B 05 BABR Repl. by LEES 12i

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(f_0(980)K^0, f_0 \rightarrow K_S^0K_S^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{332}/\Gamma$
VALUE (units $10^{-6}$ )				

**$2.7^{+1.3}_{-1.2} \pm 1.3$**  <sup>1,2</sup> LEES 12i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0K_S^0K_S^0$  decay.

$\Gamma(f_0(1710)K^0, f_0 \rightarrow K_S^0K_S^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{333}/\Gamma$
VALUE (units $10^{-6}$ )				

**$0.50^{+0.46}_{-0.24} \pm 0.11$**  <sup>1,2</sup> LEES 12i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0K_S^0K_S^0$  decay.

$\Gamma(f_0(2010)K^0, f_0 \rightarrow K_S^0K_S^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{334}/\Gamma$
VALUE (units $10^{-6}$ )				

**$0.54^{+0.21}_{-0.20} \pm 0.52$**  <sup>1,2</sup> LEES 12i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0K_S^0K_S^0$  decay.

$\Gamma(K_S^0K_S^0K_S^0 \text{ nonresonant})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{335}/\Gamma$
VALUE (units $10^{-6}$ )				

**$13.3^{+2.2}_{-2.3} \pm 2.2$**  <sup>1,2</sup> LEES 12i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0K_S^0K_S^0$  decay.

$\Gamma(K_S^0K_S^0K_L^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{336}/\Gamma$
VALUE (units $10^{-6}$ )				

**$<16$**  90 <sup>1</sup> AUBERT,B 06r BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0K^+K^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{337}/\Gamma$
VALUE (units $10^{-6}$ )				

**$27.5 \pm 1.3 \pm 2.2$**  <sup>1</sup> AUBERT 07As BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<610$  90 ALBRECHT 91e ARG  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0\phi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{338}/\Gamma$
VALUE (units $10^{-6}$ )				

**$10.0 \pm 0.5$  OUR FIT**

**$10.0 \pm 0.5$  OUR AVERAGE**

$10.4 \pm 0.5 \pm 0.6$  <sup>1</sup> PRIM 13 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

$9.7 \pm 0.5 \pm 0.5$  <sup>1</sup> AUBERT 08Bg BABR  $e^+e^- \rightarrow \Upsilon(4S)$

$11.5^{+4.5+1.8}_{-3.7-1.7}$  <sup>1</sup> BRIERE 01 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$9.2 \pm 0.7 \pm 0.6$  <sup>1</sup> AUBERT 07D BABR Repl. by AUBERT 08Bg

$9.2 \pm 0.9 \pm 0.5$  <sup>1</sup> AUBERT,B 04w BABR Repl. by AUBERT 07D

$11.2 \pm 1.3 \pm 0.8$  <sup>1</sup> AUBERT 03v BABR Repl. by AUBERT,B 04w

$10.0^{+1.6+0.7}_{-1.5-0.8}$  <sup>1</sup> CHEN 03B BELL Repl. by PRIM 13

$8.7^{+2.5}_{-2.1} \pm 1.1$  <sup>1</sup> AUBERT 01D BABR Repl. by AUBERT 03v

$<384$  90 <sup>2</sup> ABE 00c SLD  $e^+e^- \rightarrow Z$

$<21$  90 <sup>1</sup> BERGFELD 98 CLE2

$<43$  90 ASNER 96 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

$<320$  90 ALBRECHT 91B ARG  $e^+e^- \rightarrow \Upsilon(4S)$

$<380$  90 <sup>3</sup> AVERY 89B CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

$<380$  90 <sup>4</sup> AVERY 87 CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ABE 00c assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^+} = (39.7^{+1.8}_{-2.2})\%$  and  $f_{B^0} = (10.5^{+1.8}_{-2.2})\%$ .

<sup>3</sup> AVERY 89B reports  $<4.4 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

<sup>4</sup> AVERY 87 reports  $<4.7 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K^+K^-\pi^+\pi^- \text{ nonresonant})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{339}/\Gamma$
VALUE (units $10^{-6}$ )				

**$<71.7$**  90 <sup>1,2</sup> CHIANG 10 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Measured in the range  $0.7 < m_{K\pi} < 1.7$  and corrected using PS assumption for the full  $K\pi$  mass range.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0K^-\pi^+)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{340}/\Gamma$
VALUE (units $10^{-6}$ )				

**$4.5 \pm 1.3$  OUR AVERAGE**

$2.11^{+5.63+4.85}_{-5.26-4.75}$  <sup>1,2</sup> CHIANG 10 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

$4.6 \pm 1.1 \pm 0.8$  <sup>2</sup> AUBERT 07As BABR  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Measured in the range  $0.7 < m_{K\pi} < 1.7$  and corrected using PS assumption for the full  $K\pi$  mass range. The quoted result is equivalent to the upper limit of  $<13.9 \times 10^{-6}$  at 90% CL.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0\bar{K}^*(892)^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{341}/\Gamma$
VALUE (units $10^{-6}$ )				

**$0.8 \pm 0.5$  OUR AVERAGE** Error includes scale factor of 2.2.

$0.26^{+0.33+0.10}_{-0.29-0.08}$  <sup>1,2</sup> CHIANG 10 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

$1.28^{+0.35}_{-0.30} \pm 0.11$  <sup>2</sup> AUBERT 08i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<22$  90 <sup>3</sup> GODANG 02 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

$<469$  90 <sup>4</sup> ABE 00c SLD  $e^+e^- \rightarrow Z$

<sup>1</sup> Measured in the range  $0.7 < m_{K\pi} < 1.7$  and corrected using PS assumption for the full  $K\pi$  mass range. The quoted result is equivalent to the upper limit of  $<0.8 \times 10^{-6}$  at 90% CL.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to  $1.9 \times 10^{-5}$ .

<sup>4</sup> ABE 00c assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$  and  $f_{B^0} = (10.5^{+1.8}_{-2.2})\%$ .

$\Gamma(K^+K^+\pi^-\pi^- \text{ nonresonant})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{342}/\Gamma$
VALUE (units $10^{-6}$ )				

**$<6.0$**  90 <sup>1</sup> CHIANG 10 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0K^+\pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{343}/\Gamma$
VALUE (units $10^{-6}$ )				

**$<2.2$**  90 <sup>1</sup> AUBERT 07As BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<7.6$  90 <sup>1</sup> CHIANG 10 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0K^*(892)^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{344}/\Gamma$
VALUE (units $10^{-6}$ )				

**$<0.2$**  90 <sup>1</sup> CHIANG 10 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.41$  90 <sup>1</sup> AUBERT 08i BABR  $e^+e^- \rightarrow \Upsilon(4S)$

$<37$  90 <sup>2</sup> GODANG 02 CLE2  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to  $2.9 \times 10^{-5}$ .

# Meson Particle Listings

$B^0$

$\Gamma(K^*(892)^+K^*(892)^-)/\Gamma_{\text{total}}$					$\Gamma_{345}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 2.0	90	<sup>1</sup> AUBERT	08AP BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<141	90	<sup>2</sup> GODANG	02 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to  $8.9 \times 10^{-5}$ .

$\Gamma(K_1(1400)^0\phi)/\Gamma_{\text{total}}$					$\Gamma_{346}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<5.0 $\times 10^{-3}$	90	ALBRECHT	91B ARG	$e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(\phi(K\pi)_0^{*0})/\Gamma_{\text{total}}$					$\Gamma_{347}/\Gamma$
This decay refers to the coherent sum of resonant and nonresonant $J^P = 0^+ K\pi$ components with $1.13 < m_{K\pi} < 1.53 \text{ GeV}/c^2$ .					
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
4.3 $\pm$ 0.4 OUR AVERAGE					
4.3 $\pm$ 0.4 $\pm$ 0.4		<sup>1</sup> PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
4.3 $\pm$ 0.6 $\pm$ 0.4		<sup>1</sup> AUBERT	08BG BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
5.0 $\pm$ 0.8 $\pm$ 0.3		<sup>1</sup> AUBERT	07D BABR	Repl. by AUBERT 08Bg	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\phi(K\pi)_0^{*0}(1.60 < m_{K\pi} < 2.15))/\Gamma_{\text{total}}$					$\Gamma_{348}/\Gamma$
This decay refers to the coherent sum of resonant and nonresonant $J^P = 0^+ K\pi$ components with $1.60 < m_{K\pi} < 2.15 \text{ GeV}/c^2$ .					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.7	90	<sup>1</sup> AUBERT	07Ao BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^0K^-\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{349}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<31.8	90	<sup>1,2</sup> CHIANG	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Measured in the range $0.7 < m_{K\pi} < 1.7$ and corrected using PS assumption for the full $K\pi$ mass range.					
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^0\bar{K}^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{350}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<3.3	90	<sup>1,2</sup> CHIANG	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Measured in the range $0.7 < m_{K\pi} < 1.7$ and corrected using PS assumption for the full $K\pi$ mass range.					
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^0\bar{K}_0^*(1430)^0)/\Gamma_{\text{total}}$					$\Gamma_{351}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<8.4	90	<sup>1,2</sup> CHIANG	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Measured in the range $0.7 < m_{K\pi} < 1.7$ and corrected using PS assumption for the full $K\pi$ mass range.					
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^0\phi)/\Gamma_{\text{total}}$					$\Gamma_{352}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
3.9 $\pm$ 0.5 $\pm$ 0.6		<sup>1</sup> AUBERT	08BG BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.6 $\pm$ 0.7 $\pm$ 0.6		<sup>1</sup> AUBERT	07D BABR	Repl. by AUBERT 08Bg	
seen		<sup>2</sup> AUBERT,B	04W BABR	Repl. by AUBERT 07D	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Observed $181 \pm 17$ events with statistical significance greater than $10 \sigma$ .					

$\Gamma(K_0^*(1430)^0K^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{353}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.7	90	<sup>1</sup> CHIANG	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_0^*(1430)^0K_0^*(1430)^0)/\Gamma_{\text{total}}$					$\Gamma_{354}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<4.7	90	<sup>1</sup> CHIANG	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(1680)^0\phi)/\Gamma_{\text{total}}$					$\Gamma_{355}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<3.5	90	<sup>1</sup> AUBERT	07Ao BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(1780)^0\phi)/\Gamma_{\text{total}}$					$\Gamma_{356}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<2.7	90	<sup>1</sup> AUBERT	07Ao BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K^*(2045)^0\phi)/\Gamma_{\text{total}}$					$\Gamma_{357}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<15.3	90	<sup>1</sup> AUBERT	07Ao BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(K_2^*(1430)^0\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{358}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.1 $\times 10^3$	90	ALBRECHT	91B ARG	$e^+e^- \rightarrow \Upsilon(4S)$	

$\Gamma(K_2^*(1430)^0\phi)/\Gamma_{\text{total}}$					$\Gamma_{359}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
6.8 $\pm$ 0.9 OUR AVERAGE		Error includes scale factor of 1.2.			
5.5 $^{+0.9}_{-0.7} \pm 1.0$		<sup>1</sup> PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
7.5 $\pm 0.9 \pm 0.5$		<sup>1</sup> AUBERT	08BG BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
7.8 $\pm 1.1 \pm 0.6$		<sup>1</sup> AUBERT	07D BABR	Repl. by AUBERT 08Bg	
seen		<sup>2</sup> AUBERT,B	04W BABR	Repl. by AUBERT 07D	
<1400	90	ALBRECHT	91B ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> The angular distribution of $B \rightarrow \phi K^*(1430)$ provides evidence with statistical significance of $3.2 \sigma$ .					

$\Gamma(K^0\phi\phi)/\Gamma_{\text{total}}$					$\Gamma_{360}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
4.5 $\pm 0.8 \pm 0.3$		<sup>1</sup> LEES	11A BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.1 $^{+1.7}_{-1.4} \pm 0.4$		<sup>1</sup> AUBERT,BE	06H BABR	Repl. by LEES 11A	
<sup>1</sup> Assumes equal production of $B^0$ and $B^+$ at the $\Upsilon(4S)$ and for a $\phi\phi$ invariant mass below $2.85 \text{ GeV}/c^2$ .					

$\Gamma(\eta'\eta'K^0)/\Gamma_{\text{total}}$					$\Gamma_{361}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<31	90	<sup>1</sup> AUBERT,B	06P BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\eta K^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{362}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
7.6 $\pm 1.8$ OUR AVERAGE					
7.1 $^{+2.1}_{-2.0} \pm 0.4$		<sup>1,2</sup> AUBERT	09 BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
8.7 $^{+3.1}_{-2.7} \pm 1.9$		<sup>2,3</sup> NISHIDA	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
11.3 $^{+2.8}_{-1.6} \pm 0.6$		<sup>1,2</sup> AUBERT,B	06M BABR	Repl. by AUBERT 09	
<sup>1</sup> $m_{\eta K} < 3.25 \text{ GeV}/c^2$ .					
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>3</sup> $m_{\eta K} < 2.4 \text{ GeV}/c^2$					

$\Gamma(\eta'K^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{363}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<6.4	90	<sup>1,2</sup> WEDD	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<6.6	90	<sup>1,3</sup> AUBERT,B	06M BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> $m_{\eta'K} < 3.4 \text{ GeV}/c^2$ .					
<sup>3</sup> $m_{\eta'K} < 3.25 \text{ GeV}/c^2$ .					

$\Gamma(K^0\phi\gamma)/\Gamma_{\text{total}}$					$\Gamma_{364}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
2.74 $\pm 0.60 \pm 0.32$		<sup>1</sup> SAHOO	11A BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<2.7	90	<sup>1</sup> AUBERT	07q BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<8.3	90	<sup>1</sup> DRUTSKOY	04 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at $\Upsilon(4S)$ .					

$\Gamma(K^+\pi^-\gamma)/\Gamma_{\text{total}}$					$\Gamma_{365}/\Gamma$
VALUE		DOCUMENT ID	TECN	COMMENT	
(4.6 $^{+1.3}_{-1.2} \pm 0.5$ ) $\times 10^{-6}$		<sup>1,2</sup> NISHIDA	02 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> $1.25 \text{ GeV}/c^2 < M_{K\pi} < 1.6 \text{ GeV}/c^2$					

$\Gamma(K^*(892)^0\gamma)/\Gamma_{\text{total}}$					$\Gamma_{366}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
43.3 $\pm 1.5$ OUR AVERAGE					
44.7 $\pm 1.0 \pm 1.6$		<sup>1</sup> AUBERT	09Ao BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
40.1 $\pm 2.1 \pm 1.7$		<sup>2</sup> NAKAO	04 BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
45.5 $^{+7.2}_{-6.8} \pm 3.4$		<sup>3</sup> COAN	00 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	



# Meson Particle Listings

## $B^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$39.2 \pm 2.0 \pm 2.4$	90	<sup>4</sup> AUBERT,BE	04A	BABR	Repl. by AUBERT 09A0
< 110	90	ACOSTA	02G	CDF	$p\bar{p}$ at 1.8 TeV
$42.3 \pm 4.0 \pm 2.2$	90	<sup>2</sup> AUBERT	02C	BABR	Repl. by AUBERT,BE 04A
< 210	90	<sup>5</sup> ADAM	96D	DLPH	$e^+e^- \rightarrow Z$
$40 \pm 17 \pm 8$	90	<sup>6</sup> AMMAR	93	CLE2	Repl. by COAN 00
< 420	90	ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
< 240	90	<sup>7</sup> AVERY	89B	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
< 2100	90	AVERY	87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Uses  $B(\Upsilon(4S) \rightarrow B^+B^-) = (51.6 \pm 0.6)\%$  and  $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (48.4 \pm 0.6)\%$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . No evidence for a nonresonant  $K\pi\gamma$  contamination was seen; the central value assumes no contamination.  
<sup>4</sup> Uses the production ratio of charged and neutral B from  $\Upsilon(4S)$  decays  $R^{+/-} = 1.006 \pm 0.048$ .  
<sup>5</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ .  
<sup>6</sup> AMMAR 93 observed  $6.6 \pm 2.8$  events above background.  
<sup>7</sup> AVERY 89B reports  $< 2.8 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K^*(1410)\gamma)/\Gamma_{\text{total}}$		$\Gamma_{367}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 1.3 \times 10^{-4}$	90	<sup>1</sup> NISHIDA	02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^+\pi^-\gamma \text{ nonresonant})/\Gamma_{\text{total}}$		$\Gamma_{368}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 2.6 \times 10^{-6}$	90	<sup>1,2</sup> NISHIDA	02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup>  $1.25 \text{ GeV}/c^2 < M_{K\pi} < 1.6 \text{ GeV}/c^2$

$\Gamma(K^*(892)^0 X(214), X \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{369}/\Gamma$
$X(214)$ is a hypothetical particle of mass 214 MeV/ $c^2$ reported by the HyperCP experiment (PARK 05)					
VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.26</b>	90	<sup>1,2</sup> HYUN	10	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Based on scalar nature of  $X$  particle. With a vector  $X$  assumption, the upper limit is  $2.27 \times 10^{-8}$ .

$\Gamma(K^0\pi^+\pi^-\gamma)/\Gamma_{\text{total}}$		$\Gamma_{370}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.95 \pm 0.22</math> OUR AVERAGE</b>					
$1.85 \pm 0.21 \pm 0.12$		<sup>1,2</sup> AUBERT	07R	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$2.40 \pm 0.4 \pm 0.3$		<sup>2,3</sup> YANG	05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>  $M_{K\pi\pi} < 1.8 \text{ GeV}/c^2$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup>  $M_{K\pi\pi} < 2.0 \text{ GeV}/c^2$ .

$\Gamma(K^+\pi^-\pi^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{371}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>4.07 \pm 0.22 \pm 0.31</math></b>		<sup>1,2</sup> AUBERT	07R	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>  $M_{K\pi\pi} < 1.8 \text{ GeV}/c^2$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K_1(1270)^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{372}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 5.8</b>	90	<sup>1</sup> YANG	05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 700	90	<sup>2</sup> ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 89G reports  $< 0.0078$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K_1(1400)^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{373}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 1.2</b>	90	<sup>1</sup> YANG	05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 430	90	<sup>2</sup> ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 89G reports  $< 0.0048$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K_2^*(1430)^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{374}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.24 \pm 0.24</math> OUR AVERAGE</b>					
$1.22 \pm 0.25 \pm 0.10$		<sup>1</sup> AUBERT,B	04U	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.3 \pm 0.5 \pm 0.1$		<sup>1</sup> NISHIDA	02	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 40	90	<sup>2</sup> ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> ALBRECHT 89G reports  $< 4.4 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K^*(1680)^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{375}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 0.0020</b>	90	<sup>1</sup> ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 89G reports  $< 0.0022$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K_3^*(1780)^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{376}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 83</b>	90	<sup>1,2</sup> NISHIDA	05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 10000	90	<sup>3</sup> ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $B(K_3^*(1780) \rightarrow \eta K) = 0.11^{+0.05}_{-0.04}$ .  
<sup>3</sup> ALBRECHT 89G reports  $< 0.011$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(K_4^*(2045)^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{377}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 0.0043</b>	90	<sup>1</sup> ALBRECHT	89G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 89G reports  $< 0.0048$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(\rho^0\gamma)/\Gamma_{\text{total}}$		$\Gamma_{378}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.86 \pm 0.15</math> OUR AVERAGE</b>					
$0.97^{+0.24}_{-0.22} \pm 0.06$		<sup>1</sup> AUBERT	08BH	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.78^{+0.17+0.09}_{-0.16-0.10}$		<sup>1</sup> TANIGUCHI	08	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.79^{+0.22}_{-0.20} \pm 0.06$		<sup>1</sup> AUBERT	07L	BABR	Repl. by AUBERT 08BH
$1.25^{+0.37+0.07}_{-0.33-0.06}$		<sup>1</sup> MOHAPATRA	06	BELL	Repl. by TANIGUCHI 08
$0.0 \pm 0.2 \pm 0.1$	90	<sup>1</sup> AUBERT	05	BABR	Repl. by AUBERT 07L
< 0.8	90	<sup>1</sup> MOHAPATRA	05	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 1.2	90	<sup>1</sup> AUBERT	04C	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 17	90	<sup>1</sup> COAN	00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\rho^0 X(214), X \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{379}/\Gamma$
$X(214)$ is a hypothetical particle of mass 214 MeV/ $c^2$ reported by the HyperCP experiment (PARK 05)					
VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.73</b>	90	<sup>1,2</sup> HYUN	10	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> The result is the same for a scalar or vector  $X$  particle.

$\Gamma(\rho^0\gamma)/\Gamma(K^*(892)^0\gamma)$		$\Gamma_{378}/\Gamma_{366}$			
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.06^{+0.45+0.14}_{-0.43-0.16}</math></b>		TANIGUCHI	08	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\omega\gamma)/\Gamma_{\text{total}}$		$\Gamma_{380}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.44^{+0.18}_{-0.16}</math> OUR AVERAGE</b>					
$0.50^{+0.27}_{-0.23} \pm 0.09$		<sup>1</sup> AUBERT	08BH	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.40^{+0.19}_{-0.17} \pm 0.13$		<sup>1</sup> TANIGUCHI	08	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.40^{+0.24}_{-0.20} \pm 0.05$		<sup>1</sup> AUBERT	07L	BABR	Repl. by AUBERT 08BH
$0.56^{+0.34+0.05}_{-0.27-0.10}$		<sup>1</sup> MOHAPATRA	06	BELL	Repl. by TANIGUCHI 08
< 1.0	90	<sup>1</sup> AUBERT	05	BABR	Repl. by AUBERT 07L
< 0.8	90	<sup>1</sup> MOHAPATRA	05	BELL	Repl. by MOHAPATRA 06
< 1.0	90	<sup>1</sup> AUBERT	04C	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 9.2	90	<sup>1</sup> COAN	00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\phi\gamma)/\Gamma_{\text{total}}$		$\Gamma_{381}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 8.5 \times 10^{-7}</math></b>	90	<sup>1</sup> AUBERT,BE	05C	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< $0.33 \times 10^{-5}$	90	<sup>1</sup> COAN	00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

# Meson Particle Listings

$B^0$

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{382}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>5.12±0.19 OUR FIT</b>				
<b>5.13±0.24 OUR AVERAGE</b>				
5.04±0.21±0.18		<sup>1</sup> DUH	13	BELL $e^+e^- \rightarrow \gamma(4S)$
5.5 ±0.4 ±0.3		<sup>1</sup> AUBERT	07B	BABR $e^+e^- \rightarrow \gamma(4S)$
4.5 +1.4 +0.5 -1.2 -0.4		<sup>1</sup> BORNHEIM	03	CLE2 $e^+e^- \rightarrow \gamma(4S)$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
5.1 ±0.2 ±0.2		<sup>1</sup> LIN	07A	BELL Repl. by DUH 13
4.4 ±0.6 ±0.3		<sup>1</sup> CHAO	04	BELL Repl. by LIN 07A
4.7 ±0.6 ±0.2		<sup>1</sup> AUBERT	02Q	BABR Repl. by AUBERT 07B
5.4 ±1.2 ±0.5		<sup>1</sup> CASEY	02	BELL Repl. by CHAO 04
5.6 +2.3 +0.4 -2.0 -0.5		<sup>1</sup> ABE	01H	BELL Repl. by CASEY 02
4.1 ±1.0 ±0.7		<sup>1</sup> AUBERT	01E	BABR Repl. by AUBERT 02Q
< 67	90	<sup>2</sup> ABE	00C	SLD $e^+e^- \rightarrow Z$
4.3 +1.6 ±0.5 -1.4		<sup>1</sup> CRONIN-HEN..00	CLE2	Repl. by BORNHEIM 03
< 15	90	GODANG	98	CLE2 Repl. by CRONIN-HENNESSY 00
< 45	90	<sup>3</sup> ADAM	96D	DLPH $e^+e^- \rightarrow Z$
< 20	90	ASNER	96	CLE2 Repl. by GODANG 98
< 41	90	<sup>4</sup> BUSKULIC	96V	ALEP $e^+e^- \rightarrow Z$
< 55	90	<sup>5</sup> ABREU	95N	DLPH Sup. by ADAM 96D
< 47	90	<sup>6</sup> AKERS	94L	OPAL $e^+e^- \rightarrow Z$
< 29	90	<sup>1</sup> BATTLE	93	CLE2 $e^+e^- \rightarrow \gamma(4S)$
<130	90	<sup>1</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$
< 77	90	<sup>7</sup> BORTOLETTO	89	CLEO $e^+e^- \rightarrow \gamma(4S)$
<260	90	<sup>7</sup> BEBEK	87	CLEO $e^+e^- \rightarrow \gamma(4S)$
<500	90	GILES	84	CLEO $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .				
<sup>2</sup> ABE 00c assumes $B(Z \rightarrow b\bar{b})=(21.7 \pm 0.1)\%$ and the $B$ fractions $f_{B^0}=f_{B^+}= (39.7^{+1.8}_{-2.2}\%)$ and $f_{B_s}=(10.5^{+1.8}_{-2.2}\%)$ .				
<sup>3</sup> ADAM 96D assumes $f_{B^0}=f_{B^-}=0.39$ and $f_{B_s}=0.12$ .				
<sup>4</sup> BUSKULIC 96V assumes PDG 96 production fractions for $B^0, B^+, B_s, b$ baryons.				
<sup>5</sup> Assumes a $B^0, B^-$ production fraction of 0.39 and a $B_s$ production fraction of 0.12.				
<sup>6</sup> Assumes $B(Z \rightarrow b\bar{b})=0.217$ and $B_d^0(B_s^0)$ fraction 39.5% (12%).				
<sup>7</sup> Paper assumes the $\gamma(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				

$\Gamma(\pi^+\pi^-)/\Gamma(K^+\pi^-)$				$\Gamma_{382}/\Gamma_{252}$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.261±0.010 OUR FIT</b>				
<b>0.261±0.015 OUR AVERAGE</b>				
0.262±0.009±0.017	AAIJ	12AR	LHCB	$p\bar{p}$ at 7 TeV
0.259±0.017±0.016	AALTONEN	11N	CDF	$p\bar{p}$ at 1.96 TeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.21 ±0.05 ±0.03	ABULENCIA,A	06D	CDF	Repl. by AALTONEN 11N

$\Gamma(\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{383}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>1.91±0.22 OUR AVERAGE</b>					
1.83±0.21±0.13		<sup>1</sup> LEES	13D	BABR	$e^+e^- \rightarrow \gamma(4S)$
2.3 $\begin{smallmatrix} +0.4 \\ -0.5 \end{smallmatrix}$ $\begin{smallmatrix} +0.2 \\ -0.3 \end{smallmatrix}$		<sup>1</sup> CHAO	05	BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.47±0.25±0.12		<sup>1</sup> AUBERT	07Bc	BABR	Repl. by LEES 13D
1.17±0.32±0.10		<sup>1</sup> AUBERT	05L	BABR	Repl. by AUBERT 07Bc
< 3.6	90	<sup>1</sup> AUBERT	03L	BABR	$e^+e^- \rightarrow \gamma(4S)$
2.1 ±0.6 ±0.3		<sup>1</sup> AUBERT	03s	BABR	Repl. by AUBERT 05L
< 4.4	90	<sup>1</sup> BORNHEIM	03	CLE2	$e^+e^- \rightarrow \gamma(4S)$
1.7 ±0.6 ±0.2		<sup>1</sup> LEE	03	BELL	Repl. by CHAO 05
< 5.7	90	<sup>1</sup> ASNER	02	CLE2	$e^+e^- \rightarrow \gamma(4S)$
< 6.4	90	<sup>1</sup> CASEY	02	BELL	$e^+e^- \rightarrow \gamma(4S)$
< 9.3	90	GODANG	98	CLE2	Repl. by ASNER 02
< 9.1	90	ASNER	96	CLE2	Repl. by GODANG 98
<60	90	<sup>2</sup> ACCIARRI	95H	L3	$e^+e^- \rightarrow Z$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> ACCIARRI 95H assumes $f_{B^0} = 39.5 \pm 4.0$ and $f_{B_s} = 12.0 \pm 3.0\%$ .					

$\Gamma(\eta\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{384}/\Gamma$				
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT					
<b>0.41±0.17±0.05 0.15-0.07</b>									
• • • We do not use the following data for averages, fits, limits, etc. • • •									
< 1.5	90	<sup>2</sup> AUBERT	08AH	BABR	$e^+e^- \rightarrow \gamma(4S)$				
< 1.3	90	<sup>2</sup> AUBERT	06W	BABR	Repl. by AUBERT 08AH				
< 2.5	90	<sup>2</sup> CHANG	05A	BELL	Repl. by PAL 15				
< 2.5	90	<sup>2</sup> AUBERT,B	04D	BABR	Repl. by AUBERT 06W				
< 2.9	90	<sup>2</sup> RICHICHI	00	CLE2	$e^+e^- \rightarrow \gamma(4S)$				
< 8	90	BEHRENS	98	CLE2	Repl. by RICHICHI 00				
< 250	90	<sup>3</sup> ACCIARRI	95H	L3	$e^+e^- \rightarrow Z$				
<1800	90	<sup>2</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$				

<sup>1</sup> PAL 15 signal significance is 3.0 standard deviations. The measurement corresponds to 90% CL upper limit of  $< 6.5 \times 10^{-7}$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>3</sup> ACCIARRI 95H assumes  $f_{B^0}=39.5 \pm 4.0$  and  $f_{B_s}=12.0 \pm 3.0\%$ .

$\Gamma(\eta\eta)/\Gamma_{\text{total}}$					$\Gamma_{385}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 1.0	90	<sup>1</sup> AUBERT	09AV	BABR	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 1.8	90	<sup>1</sup> AUBERT,B	06V	BABR	Repl. by AUBERT 09AV
< 2.0	90	<sup>1</sup> CHANG	05A	BELL	$e^+e^- \rightarrow \gamma(4S)$
< 2.8	90	<sup>1</sup> AUBERT,B	04X	BABR	$e^+e^- \rightarrow \gamma(4S)$
< 18	90	BEHRENS	98	CLE2	$e^+e^- \rightarrow \gamma(4S)$
<410	90	<sup>2</sup> ACCIARRI	95H	L3	$e^+e^- \rightarrow Z$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .					
<sup>2</sup> ACCIARRI 95H assumes $f_{B^0} = 39.5 \pm 4.0$ and $f_{B_s} = 12.0 \pm 3.0\%$ .					

$\Gamma(\eta'\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{386}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>1.2±0.6 OUR AVERAGE</b>		Error includes scale factor of 1.7.				
0.9±0.4±0.1		<sup>1</sup> AUBERT	08AH	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
2.8±1.0±0.3		<sup>1</sup> SCHUEMANN	06	BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.8+0.8 -0.6±0.1		<sup>1</sup> AUBERT	06W	BABR	Repl. by AUBERT 08AH	
1.0+1.4 -1.0±0.8	90	<sup>1</sup> AUBERT,B	04D	BABR	Repl. by AUBERT 06W	
< 5.7	90	<sup>1</sup> RICHICHI	00	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
<11	90	BEHRENS	98	CLE2	Repl. by RICHICHI 00	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .						

$\Gamma(\eta'\eta')/\Gamma_{\text{total}}$					$\Gamma_{387}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
$< 1.7$	90	<sup>1</sup> AUBERT	09AV	BABR	$e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$< 6.5$	90	<sup>1</sup> SCHUEMANN	07	BELL	$e^+e^- \rightarrow \gamma(4S)$	
$< 2.4$	90	<sup>1</sup> AUBERT,B	06V	BABR	Repl. by AUBERT 09AV	
$<10$	90	<sup>1</sup> AUBERT,B	04X	BABR	Repl. by AUBERT,B 06V	
$<47$	90	BEHRENS	98	CLE2	$e^+e^- \rightarrow \gamma(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .						

$\Gamma(\eta'\eta)/\Gamma_{\text{total}}$
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See key on page 601

## Meson Particle Listings

 $B^0$  $\Gamma(\eta f_0(980), f_0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{392}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.4</b>	90	<sup>1</sup> AUBERT	07Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\omega \eta)/\Gamma_{\text{total}}$   $\Gamma_{393}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.94 \pm 0.35</math> <math>-0.30 \pm 0.09</math></b>		<sup>1</sup> AUBERT	09AV BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.9	90	<sup>1</sup> AUBERT,B	05K BABR	Repl. by AUBERT 09AV
$4.0 \pm 1.3$ $-1.2 \pm 0.4$		<sup>1</sup> AUBERT,B	04X BABR	Repl. by AUBERT,B 05K
<12	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\omega \eta')/\Gamma_{\text{total}}$   $\Gamma_{394}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.01 \pm 0.46</math> <math>-0.38 \pm 0.09</math></b>		<sup>1</sup> AUBERT	09AV BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 2.2	90	<sup>1</sup> SCHUEMANN 07	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
< 2.8	90	<sup>1</sup> AUBERT,B	04X BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<60	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\omega \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{395}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.6</b>	90	<sup>1</sup> AUBERT	09H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.5	90	<sup>1</sup> AUBERT,B	06T BABR	Repl. by AUBERT 09H
< 3.3	90	<sup>1</sup> AUBERT	05O BABR	Repl. by AUBERT,B 06T
<11	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\omega f_0(980), f_0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{396}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.5</b>	90	<sup>1</sup> AUBERT	09H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.5	90	<sup>1</sup> AUBERT,B	06T BABR	Repl. by AUBERT 09H

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\omega \omega)/\Gamma_{\text{total}}$   $\Gamma_{397}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.2 \pm 0.3</math> <math>-0.2</math></b>		<sup>1</sup> LEES	14 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 4.0	90	<sup>1</sup> AUBERT,B	06T BABR	Repl. by LEES 14
<19	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\phi \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{398}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.15</b>	90	<sup>1</sup> KIM	12A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.28	90	<sup>1</sup> AUBERT,B	06C BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<1.0	90	<sup>1</sup> AUBERT,B	04D BABR	Repl. by AUBERT,B 06C
<5	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\phi \eta)/\Gamma_{\text{total}}$   $\Gamma_{399}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.5</b>	90	<sup>1</sup> AUBERT	09AV BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.6	90	<sup>1</sup> AUBERT,B	06V BABR	Repl. by AUBERT 09AV
<1.0	90	<sup>1</sup> AUBERT,B	04X BABR	Repl. by AUBERT,B 06V
<9	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\phi \eta')/\Gamma_{\text{total}}$   $\Gamma_{400}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.5</b>	90	<sup>1</sup> SCHUEMANN 07	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.1	90	<sup>1</sup> AUBERT	09AV BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
< 1.0	90	<sup>1</sup> AUBERT,B	06V BABR	Repl. by AUBERT 09AV
< 4.5	90	<sup>1</sup> AUBERT,B	04X BABR	Repl. by AUBERT,B 06V
<31	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\phi \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{401}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.33</b>	90	<sup>1</sup> AUBERT	08BK BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<156	90	<sup>2</sup> ABE	00C SLD	$e^+ e^- \rightarrow Z$
< 13	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> ABE 00C assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0} = f_{B^+} = (39.7 \pm 1.8)\%$  and  $f_{B_s} = (10.5 \pm 1.8)\%$ . $\Gamma(\phi f_0(980), f_0 \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{402}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.38</b>	90	<sup>1</sup> AUBERT	08BK BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\phi \omega)/\Gamma_{\text{total}}$   $\Gamma_{403}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.7</b>	90	<sup>1</sup> LEES	14 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.2	90	<sup>1</sup> AUBERT,B	06T BABR	Repl. by LEES 14
<21	90	<sup>1</sup> BERGFELD	98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\phi \phi)/\Gamma_{\text{total}}$   $\Gamma_{404}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.8 \times 10^{-8}</math></b>	90	AAIJ	15As LHCb	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2 $\times 10^{-7}$	90	<sup>1</sup> AUBERT	08BK BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<1.5 $\times 10^{-6}$	90	<sup>1</sup> AUBERT,B	04X BABR	Repl. by AUBERT 08BK
<3.21 $\times 10^{-4}$	90	<sup>2</sup> ABE	00C SLD	$e^+ e^- \rightarrow Z$
<1.2 $\times 10^{-5}$	90	<sup>1</sup> BERGFELD	98 CLE2	
<3.9 $\times 10^{-5}$	90	ASNER	96 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> ABE 00C assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0} = f_{B^+} = (39.7 \pm 1.8)\%$  and  $f_{B_s} = (10.5 \pm 1.8)\%$ . $\Gamma(a_0(980) \pm \pi^\mp, a_0^\pm \rightarrow \eta \pi^\pm)/\Gamma_{\text{total}}$   $\Gamma_{405}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.1</b>	90	<sup>1</sup> AUBERT	07Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<5.1	90	<sup>1</sup> AUBERT,BE	04 BABR	Repl. by AUBERT 07Y

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(a_0(1450) \pm \pi^\mp, a_0^\pm \rightarrow \eta \pi^\pm)/\Gamma_{\text{total}}$   $\Gamma_{406}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.3</b>	90	<sup>1</sup> AUBERT	07Y BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ . $\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{407}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.2 <math>\times 10^{-4}</math></b>	90	<sup>1</sup> ALBRECHT	90B ARG	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 90B limit assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at  $\Upsilon(4S)$ . $\Gamma(\rho^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{408}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.0 <math>\pm 0.5</math> OUR AVERAGE</b>				
3.0 $\pm 0.5 \pm 0.7$		<sup>1,2</sup> KUSAKA	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
1.4 $\pm 0.6 \pm 0.3$		<sup>1</sup> AUBERT	04Z BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
1.6 $\pm 2.0$ $-1.4 \pm 0.8$		<sup>1</sup> JESSOP	00 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.12 $\pm 0.88 \pm 0.60$ $-0.82 \pm 0.76$		<sup>1</sup> DRAGIC	06 BELL	Repl. by KUSAKA 08
5.1 $\pm 1.6 \pm 0.9$		DRAGIC	04 BELL	Repl. by DRAGIC 06
< 5.3	90	<sup>1</sup> GORDON	02 BELL	Repl. by DRAGIC 04
< 24	90	ASNER	96 CLEO	Repl. by JESSOP 00
<400	90	<sup>1</sup> ALBRECHT	90B ARG	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> This is the first measurement that excludes contributions from  $\rho(1450)$  and  $\rho(1570)$  resonances. $\Gamma(\rho^\mp \pi^\pm)/\Gamma_{\text{total}}$   $\Gamma_{409}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>23.0 <math>\pm 2.3</math> OUR AVERAGE</b>				
22.6 $\pm 1.1 \pm 4.4$		<sup>1,2</sup> KUSAKA	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
22.6 $\pm 1.8 \pm 2.2$		<sup>1</sup> AUBERT	03T BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
27.6 $\pm 8.4$ $-7.4 \pm 4.2$		<sup>1</sup> JESSOP	00 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

# Meson Particle Listings

$B^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$20.8^{+6.0+2.8}_{-6.3-3.1}$		<sup>1</sup> GORDON	02	BELL	Repl. by KUSAKA 08
< 88	90	ASNER	96	CLE2	Repl. by JESSOP 00
< 520	90	<sup>1</sup> ALBRECHT	90B	ARG	$e^+e^- \rightarrow \gamma(4S)$
< 5200	90	<sup>3</sup> BEBEK	87	CLEO	$e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>2</sup> This is the first measurement that excludes contributions from  $\rho(1450)$  and  $\rho(1570)$  resonances.  
<sup>3</sup> BEBEK 87 reports  $< 6.1 \times 10^{-3}$  assuming the  $\gamma(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

## $\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{410}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;11.2 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> VANHOEFER	14	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<23.1 $\times 10^{-6}$	90	<sup>1</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \gamma(4S)$
<19.3 $\times 10^{-6}$	90	<sup>1</sup> CHIANG	08	BELL Repl. by VANHOEFER 14
< 2.3 $\times 10^{-4}$	90	<sup>2</sup> ADAM	96D	DLPH $e^+e^- \rightarrow Z$
< 2.8 $\times 10^{-4}$	90	<sup>3</sup> ABREU	95N	DLPH Sup. by ADAM 96D
< 6.7 $\times 10^{-4}$	90	<sup>1</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>2</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ .  
<sup>3</sup> Assumes a  $B^0$ ,  $B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12.

## $\Gamma(\rho^0\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{411}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 8.8</b>	90	<sup>1</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<12.0	90	<sup>1</sup> VANHOEFER	14	BELL $e^+e^- \rightarrow \gamma(4S)$
<12.0	90	<sup>1</sup> CHIANG	08	BELL Repl. by VANHOEFER 14

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(\rho^0\rho^0)/\Gamma_{\text{total}}$ $\Gamma_{412}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.96<math>\pm</math>0.15 OUR FIT</b>				
<b>0.97<math>\pm</math>0.24 OUR AVERAGE</b>				
1.02 $\pm$ 0.30 $\pm$ 0.15		<sup>1,2</sup> VANHOEFER	14	BELL $e^+e^- \rightarrow \gamma(4S)$
0.92 $\pm$ 0.32 $\pm$ 0.14		<sup>2</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.4 $\pm$ 0.4 $\pm$ $^{+0.2}_{-0.3}$		<sup>2</sup> CHIANG	08	BELL Repl. by VANHOEFER 14
1.07 $\pm$ 0.33 $\pm$ 0.19		<sup>2</sup> AUBERT	07G	BABR Repl. by AUBERT 08B
< 1.1	90	<sup>2</sup> AUBERT	05I	BABR Repl. by AUBERT 07G
< 2.1	90	<sup>2</sup> AUBERT	03V	BABR Repl. by AUBERT 05I
< 18	90	<sup>3</sup> GODANG	02	CLE2 $e^+e^- \rightarrow \gamma(4S)$
<136	90	<sup>4</sup> ABE	00C	SLD $e^+e^- \rightarrow Z$
<280	90	<sup>2</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$
<290	90	<sup>5</sup> BORTOLETTO	089	CLEO $e^+e^- \rightarrow \gamma(4S)$
<430	90	<sup>5</sup> BEBEK	87	CLEO $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Signal significance 3.4 standard deviations.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>3</sup> Assumes a helicity 00 configuration. For a helicity 11 configuration, the limit decreases to  $1.4 \times 10^{-5}$ .  
<sup>4</sup> ABE 00C assumes  $B(Z \rightarrow b\bar{b}) = (21.7 \pm 0.1)\%$  and the  $B$  fractions  $f_{B^0} = f_{B^+} = (39.7^{+1.8}_{-2.2})\%$  and  $f_{B_s} = (10.5^{+1.8}_{-2.2})\%$ .  
<sup>5</sup> Paper assumes the  $\gamma(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

## $\Gamma(\rho^0\rho^0)/\Gamma(K^*(892)^0\phi)$ $\Gamma_{412}/\Gamma_{338}$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>9.5<math>\pm</math>1.5 OUR FIT</b>			
<b>9.4<math>\pm</math>1.7<math>\pm</math>0.9</b>	AAIJ	15T	LHCB $pp$ at 7, 8 TeV

## $\Gamma(f_0(980)\pi^+\pi^-, f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{413}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.0 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> VANHOEFER	14	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.8 $\times 10^{-6}$	90	<sup>1</sup> CHIANG	08	BELL $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(\rho^0 f_0(980), f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{414}/\Gamma$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>7.8<math>\pm</math>2.2<math>\pm</math>1.1</b>		<sup>1,2</sup> VANHOEFER	14	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<8.1	90	AAIJ	15T	LHCB $pp$ at 7, 8 TeV
<4.0	90	<sup>2</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \gamma(4S)$
<3	90	<sup>2</sup> CHIANG	08	BELL Repl. by VANHOEFER 14
<5.3	90	<sup>2</sup> AUBERT	07G	BABR Repl. by AUBERT 08B

- <sup>1</sup> Signal significance of 3.1 standard deviations.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(f_0(980)f_0(980), f_0 \rightarrow \pi^+\pi^-, f_0 \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{415}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.19</b>	90	<sup>1</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.2	90	<sup>1</sup> VANHOEFER	14	BELL $e^+e^- \rightarrow \gamma(4S)$
<0.1	90	<sup>1</sup> CHIANG	08	BELL Repl. by VANHOEFER 14
<0.16	90	<sup>1</sup> AUBERT	07G	BABR Repl. by AUBERT 08B

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(f_0(980)f_0(980), f_0 \rightarrow \pi^+\pi^-, f_0 \rightarrow K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_{416}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.23</b>	90	<sup>1</sup> AUBERT	08B	BABR $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

## $\Gamma(a_1(1260)\mp\pi^\pm)/\Gamma_{\text{total}}$ $\Gamma_{417}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>26 <math>\pm</math> 5 OUR AVERAGE</b>				Error includes scale factor of 1.9.
22.2 $\pm$ 2.0 $\pm$ 2.8		<sup>1,2</sup> DALSENO	12	BELL $e^+e^- \rightarrow \gamma(4S)$
33.2 $\pm$ 3.8 $\pm$ 3.0		<sup>2,3</sup> AUBERT	06V	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 630	90	<sup>2</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$
< 490	90	<sup>4</sup> BORTOLETTO	089	CLEO $e^+e^- \rightarrow \gamma(4S)$
<1000	90	<sup>4</sup> BEBEK	87	CLEO $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> DALSENO 12 reports  $B(B^0 \rightarrow a_1^\pm \pi^\mp) B(a_1^\pm \rightarrow \pi^\pm \pi^+ \pi^-) = (11.1 \pm 1.0 \pm 1.4) \times 10^{-6}$  which we rescaled assuming  $a_1(1260)$  decays only to  $3\pi$  and  $B(a_1^\pm \rightarrow \pi^\pm \pi^+ \pi^-) = 0.5$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>3</sup> Assumes  $a_1(1260)$  decays only to  $3\pi$  and  $B(a_1^\pm \rightarrow \pi^\pm \pi^\mp \pi^\pm) = 0.5$ .  
<sup>4</sup> Paper assumes the  $\gamma(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

## $\Gamma(a_2(1320)\mp\pi^\pm)/\Gamma_{\text{total}}$ $\Gamma_{418}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.3 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> DALSENO	12	BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0 $\times 10^{-4}$	90	<sup>2</sup> BORTOLETTO	089	CLEO $e^+e^- \rightarrow \gamma(4S)$
<1.4 $\times 10^{-3}$	90	<sup>2</sup> BEBEK	87	CLEO $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> DALSENO 12 reports  $B(B^0 \rightarrow a_2^\pm \pi^\mp) B(a_2^\pm \rightarrow \pi^\pm \pi^+ \pi^-) < 2.2 \times 10^{-6}$  which we rescaled using  $B(a_2^\pm \rightarrow \pi^\pm \pi^+ \pi^-) = 1/2 B(a_2^\pm \rightarrow 3\pi) = 0.35 \pm 0.013$ .  
<sup>2</sup> Paper assumes the  $\gamma(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

## $\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{419}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.1 <math>\times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> ALBRECHT 90B limit assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at  $\gamma(4S)$ .

## $\Gamma(\rho^+\rho^-)/\Gamma_{\text{total}}$ $\Gamma_{420}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>27.7<math>\pm</math>1.9 OUR AVERAGE</b>				
28.3 $\pm$ 1.5 $\pm$ 1.5		<sup>1</sup> VANHOEFER	16	BELL $e^+e^- \rightarrow \gamma(4S)$
25.5 $\pm$ 2.1 $\pm$ $^{+3.6}_{-3.9}$		<sup>1</sup> AUBERT	07Bf	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
22.8 $\pm$ 3.8 $\pm$ $^{+2.3}_{-2.6}$		<sup>1</sup> SOMOV	06	BELL Repl. by VANHOEFER 16
25 $^{+7}_{-6}$ $^{+5}_{-6}$		<sup>1</sup> AUBERT	04G	BABR Repl. by AUBERT, B 04R
30 $^{+4}_{-4}$ $^{+5}_{-5}$		<sup>1,2</sup> AUBERT, B	04R	BABR Repl. by AUBERT 07Bf
<2200	90	<sup>1</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>2</sup> The quoted result is obtained after combining with AUBERT 04G result by AUBERT 04R alone gives  $(33 \pm 4 \pm 5) \times 10^{-6}$ .

## $\Gamma(a_1(1260)^0\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{421}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 <math>\times 10^{-3}</math></b>	90	<sup>1</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> ALBRECHT 90B limit assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at  $\gamma(4S)$ .

## $\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{422}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.5</b>	90	<sup>1</sup> AUBERT	08A	BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 2.0	90	<sup>1</sup> JEN	06	BELL $e^+e^- \rightarrow \gamma(4S)$
< 1.2	90	<sup>1</sup> AUBERT, B	04D	BABR Repl. by AUBERT 08A
< 1.9	90	<sup>1</sup> WANG	04A	BELL $e^+e^- \rightarrow \gamma(4S)$
< 3	90	<sup>1</sup> AUBERT	01G	BABR $e^+e^- \rightarrow \gamma(4S)$
< 5.5	90	<sup>1</sup> JESSOP	00	CLE2 $e^+e^- \rightarrow \gamma(4S)$
< 14	90	<sup>1</sup> BERGFELD	98	CLE2 Repl. by JESSOP 00
<460	90	<sup>2</sup> ALBRECHT	90B	ARG $e^+e^- \rightarrow \gamma(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .  
<sup>2</sup> ALBRECHT 90B limit assumes equal production of  $B^0 \bar{B}^0$  and  $B^+ B^-$  at  $\gamma(4S)$ .

See key on page 601

## Meson Particle Listings

 $B^0$ 

$\Gamma(\pi^+\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{423}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<9.0 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .					

$\Gamma(a_1(1260)^+\rho^-)/\Gamma_{\text{total}}$					$\Gamma_{424}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
$< 61$	90	<sup>1,2</sup> AUBERT, B 060	BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<3400$	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					
<sup>2</sup> Assumes $a_1(1260)$ decays only to $3\pi$ and $B(a_1^\pm \rightarrow \pi^\pm \pi^\mp \pi^\pm) = 0.5$ .					

$\Gamma(a_1(1260)^0\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{425}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.4 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .					

$\Gamma(b_1^-\pi^\pm, b_1^-\pi^\mp \rightarrow \omega\pi^\mp)/\Gamma_{\text{total}}$				$\Gamma_{426}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT	
$10.9 \pm 1.2 \pm 0.9$	<sup>1</sup> AUBERT	07BI BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(b_1^0\pi^0, b_1^0\rightarrow\omega\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{427}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.9</b>	90	<sup>1</sup> AUBERT	08AG BABR	$e^+e^-\rightarrow\Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(b_1^-\rho^+, b_1^-\omega\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{428}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.4 \times 10^{-6}$	90	<sup>1</sup> AUBERT	09AF BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(b_1^0\rho^0, b_1^0\rightarrow\omega\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{429}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.4\times 10^{-6}$	90	<sup>1</sup> AUBERT	09AF BABR	$e^+e^-\rightarrow\Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(\pi^+\pi^+\pi^+\pi^-\pi^-\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{430}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.0 \times 10^{-3}$	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .					

$\Gamma(a_1(1260)^+a_1(1260)^-, a_1^+ \rightarrow 2\pi^+\pi^-, a_1^- \rightarrow 2\pi^-\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{431}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>11.8 \pm 2.6 \pm 1.6</math></b>		<sup>1</sup> AUBERT	09AL BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<6000	90	<sup>1</sup> ALBRECHT	90B ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<2800	90	<sup>2</sup> BORTOLETTO	89 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .					
<sup>2</sup> BORTOLETTO 89 reports $<3.2 \times 10^{-3}$ assuming the $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.					

$\Gamma(\pi^+\pi^+\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{432}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.1 \times 10^{-2}$	90	<sup>1</sup> ALBRECHT 90B	ARG	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> ALBRECHT 90B limit assumes equal production of $B^0\bar{B}^0$ and $B^+B^-$ at $\Upsilon(4S)$ .					

$\Gamma(p\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{433}/\Gamma$
VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$1.47^{+0.62+0.35}_{-0.51-0.14}$		<sup>1</sup> AAIJ	13Bq LHCb	$p\bar{p}$ at 7 TeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 11	90	<sup>2</sup> TSAI	07 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 41	90	<sup>2</sup> CHANG	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 27	90	<sup>2</sup> AUBERT	04u BABR	$e^+e^- \rightarrow \Upsilon(4S)$
< 140	90	<sup>2</sup> BORNHEIM	03 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 120	90	<sup>2</sup> ABE	02o BELL	$e^+e^- \rightarrow \Upsilon(4S)$
< 700	90	<sup>2</sup> COAN	99 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
< 1800	90	<sup>3</sup> BUSKULIC	96v ALEP	$e^+e^- \rightarrow Z$
< 35000	90	<sup>4</sup> ABREU	95N DLPH	Sup. by ADAM 960
< 3400	90	<sup>5</sup> BORTOLETTO	89 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
< 12000	90	<sup>6</sup> ALBRECHT	88f ARG	$e^+e^- \rightarrow \Upsilon(4S)$
< 17000	90	<sup>5</sup> BEBEK	87 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Uses normalization mode  $B(B^0 \rightarrow K^+\pi^-) = (19.55 \pm 0.54) \times 10^{-6}$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> BUSKULIC 96V assumes PDG 96 production fractions for  $B^0, B^+, B_s, b$  baryons.  
<sup>4</sup> Assumes a  $B^0, B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12.  
<sup>5</sup> Paper assumes the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.  
<sup>6</sup> ALBRECHT 88F reports  $<1.3 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(p\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{434}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.5</b>	90	<sup>1</sup> BEBEK	89 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					

- <sup>1</sup> BEBEK 89 reports  $<2.9 \times 10^{-4}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0\bar{B}^0$ . We rescale to 50%.  
<sup>2</sup> Assumes a  $B^0, B^-$  production fraction of 0.39 and a  $B_s$  production fraction of 0.12.  
<sup>3</sup> ALBRECHT 88F reports  $6.0 \pm 2.0 \pm 2.2$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0\bar{B}^0$ . We rescale to 50%.

$\Gamma(p\bar{p}K^0)/\Gamma_{\text{total}}$					$\Gamma_{435}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>2.66±0.32 OUR AVERAGE</b>						
$2.51^{+0.35}_{-0.29} \pm 0.21$		<sup>1,2</sup> CHEN	08c BELL	$e^+e^- \rightarrow \Upsilon(4S)$		
$3.0 \pm 0.5 \pm 0.3$		<sup>2</sup> AUBERT	07Av BABR	$e^+e^- \rightarrow \Upsilon(4S)$		
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
$2.40^{+0.64}_{-0.44} \pm 0.28$		<sup>2,3,4</sup> WANG	05A BELL	Repl. by CHEN 08c		
$1.88^{+0.77}_{-0.60} \pm 0.23$		<sup>2,3,5</sup> WANG	04 BELL	Repl. by WANG 05A		
<7.2	90	<sup>2,3</sup> ABE	02k BELL	Repl. by WANG 04		

- <sup>1</sup> Explicitly vetoes resonant production of  $p\bar{p}$  from charmonium states.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> Explicitly vetoes resonant production of  $p\bar{p}$  from charmonium states and  $pK^0$  production from  $\Lambda_c$ .  
<sup>4</sup> Provides also results with  $M_{p\bar{p}} < 2.85 \text{ GeV}/c^2$  and angular asymmetry of  $p\bar{p}$  system.  
<sup>5</sup> The branching fraction for  $M_{p\bar{p}} < 2.85$  is also reported.

$\Gamma(\Theta(1540)^+\bar{p}, \Theta^+ \rightarrow pK_S^0)/\Gamma_{\text{total}}$					$\Gamma_{436}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.05</b>	90	<sup>1</sup> AUBERT	07Av BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.23	90	<sup>1</sup> WANG	05A BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(f_J(2220)K^0, f_J \rightarrow p\bar{p})/\Gamma_{\text{total}}$					$\Gamma_{437}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.45</b>	90	<sup>1</sup> AUBERT	07AV	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(p\bar{p}K^*(892)^0)/\Gamma_{\text{total}}$				$\Gamma_{438}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.24^{+0.28}_{-0.25}</math> OUR AVERAGE</b>				
$1.18^{+0.29}_{-0.25} \pm 0.11$		<sup>1,2</sup> CHEN	08c BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$1.47 \pm 0.45 \pm 0.40$		<sup>2</sup> AUBERT	07Av BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<7.6$	90	<sup>2</sup> WANG	04 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Explicitly vetoes resonant production of $p\bar{p}$ from charmonium states.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(f_J(2220)K_0^*, f_J \rightarrow p\bar{p})/\Gamma_{\text{total}}$					$\Gamma_{439}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.15</b>	90	<sup>1</sup> AUBERT	07AV BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .					

$\Gamma(p\bar{\Lambda}\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{440}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.14 \pm 0.29</math> OUR AVERAGE</b>					
$3.07 \pm 0.31 \pm 0.23$		<sup>1</sup> AUBERT	09Ac BABR	$e^+e^- \rightarrow \Upsilon(4S)$	
$3.23^{+0.33}_{-0.29} \pm 0.29$		<sup>1</sup> WANG	07c BELL	$e^+e^- \rightarrow \Upsilon(4S)$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$2.62^{+0.44}_{-0.40} \pm 0.31$		<sup>1,2</sup> WANG	05A BELL	Repl. by WANG 07c	
$3.97^{+1.00}_{-0.80} \pm 0.56$		<sup>1</sup> WANG	03 BELL	Repl. by WANG 05A	
$< 13$	90	<sup>1</sup> COAN	99 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$	
$< 180$	90	<sup>3</sup> ALBRECHT	88f ARG	$e^+e^- \rightarrow \Upsilon(4S)$	

# Meson Particle Listings

$B^0$

$\Gamma(\rho\Xi(1385)^-)/\Gamma_{\text{total}}$				$\Gamma_{442}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.26</b>	90	<sup>1</sup> WANG	07c	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\Delta^0\bar{\Lambda})/\Gamma_{\text{total}}$				$\Gamma_{443}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.93</b>	90	<sup>1</sup> WANG	07c	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\rho\bar{K}K^-)/\Gamma_{\text{total}}$				$\Gamma_{444}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.82</b>	90	<sup>1</sup> WANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\rho\bar{\Lambda}D^-)/\Gamma_{\text{total}}$				$\Gamma_{445}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
<b>25.1±2.6±3.5</b>		<sup>1</sup> CHANG	15	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\rho\bar{\Lambda}D^{*-})/\Gamma_{\text{total}}$				$\Gamma_{446}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
<b>33.6±6.3±4.4</b>		<sup>1</sup> CHANG	15	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\rho\Sigma^0\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{447}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.8 × 10<sup>-6</sup></b>	90	<sup>1</sup> WANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{\Lambda}\Lambda)/\Gamma_{\text{total}}$				$\Gamma_{448}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.32</b>	90	<sup>1</sup> TSAI	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.				• • •
<0.69	90	<sup>1</sup> CHANG	05	BELL Repl. by TSAI 07
<1.2	90	<sup>1</sup> BORNHEIM	03	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<1.0	90	<sup>1</sup> ABE	02o	BELL Repl. by CHANG 05
<3.9	90	<sup>1</sup> COAN	99	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{\Lambda}\Lambda K^0)/\Gamma_{\text{total}}$				$\Gamma_{449}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
<b>4.76<sup>+0.87</sup><sub>-0.68</sub>±0.61</b>		<sup>1,2</sup> CHANG	09	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Excluding charmonium events in $2.85 < m_{\Lambda\bar{\Lambda}} < 3.128$ GeV/c <sup>2</sup> and $3.315 < m_{\Lambda\bar{\Lambda}} < 3.735$ GeV/c <sup>2</sup> . Measurements in various $m_{\Lambda\bar{\Lambda}}$ bins are also reported.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{\Lambda}\Lambda K^{*0})/\Gamma_{\text{total}}$				$\Gamma_{450}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT
<b>2.46<sup>+0.87</sup><sub>-0.72</sub>±0.34</b>		<sup>1,2</sup> CHANG	09	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Excluding charmonium events in $2.85 < m_{\Lambda\bar{\Lambda}} < 3.128$ GeV/c <sup>2</sup> and $3.315 < m_{\Lambda\bar{\Lambda}} < 3.735$ GeV/c <sup>2</sup> . Measurements in various $m_{\Lambda\bar{\Lambda}}$ bins are also reported.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{\Lambda}\Lambda D^0)/\Gamma_{\text{total}}$				$\Gamma_{451}/\Gamma$
VALUE (units $10^{-5}$ )		DOCUMENT ID	TECN	COMMENT
<b>1.00<sup>+0.30</sup><sub>-0.26</sub> OUR AVERAGE</b>				
0.98 <sup>+0.29</sup> <sub>-0.26</sub> ±0.19		<sup>1,2</sup> LEES	14b	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.05 <sup>+0.57</sup> <sub>-0.44</sub> ±0.14		<sup>2</sup> CHANG	09	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Evidence for 3.4 st. dev. signal significance.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D^0\Sigma^0\bar{\Lambda} + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_{452}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.1 × 10<sup>-5</sup></b>	90	<sup>1,2</sup> LEES	14b	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Here $\Sigma^0 \rightarrow \Lambda\gamma$ .				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\Delta^0\bar{D}^0)/\Gamma_{\text{total}}$				$\Gamma_{453}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.0015</b>	90	<sup>1</sup> BORTOLETTO89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> BORTOLETTO 89 reports < 0.0018 assuming $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				

$\Gamma(D^{++}\bar{D}^{--})/\Gamma_{\text{total}}$				$\Gamma_{454}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 × 10<sup>-4</sup></b>	90	<sup>1</sup> BORTOLETTO89	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> BORTOLETTO 89 reports < 1.3 × 10 <sup>-4</sup> assuming $\Upsilon(4S)$ decays 43% to $B^0\bar{B}^0$ . We rescale to 50%.				

$\Gamma(\bar{D}^0\rho\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{455}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>1.04±0.07 OUR AVERAGE</b>				
1.02±0.04±0.06		<sup>1,2</sup> DEL-AMO-SA...12	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.18±0.15±0.16		<sup>2</sup> ABE	02w	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.				• • •
1.13±0.06±0.08		<sup>2</sup> AUBERT,B	06s	BABR Repl. by DEL-AMO-SANCHEZ 12
<sup>1</sup> Uses the values of $D$ and $D^*$ branching fractions from PDG 08.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D_s^-\bar{\Lambda}\rho)/\Gamma_{\text{total}}$				$\Gamma_{456}/\Gamma$
VALUE (units $10^{-5}$ )		DOCUMENT ID	TECN	COMMENT
<b>2.8±0.8±0.3</b>		<sup>1,2</sup> MEDVEDEVA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				
<sup>2</sup> MEDVEDEVA 07 reports $(2.9\pm0.7\pm0.5\pm0.4)\times10^{-5}$ from a measurement of $[\Gamma(B^0\rightarrow D_s^-\bar{\Lambda}\rho)/\Gamma_{\text{total}}]\times[B(D_s^+\rightarrow\phi\pi^+)]$ assuming $B(D_s^+\rightarrow\phi\pi^+)=(4.4\pm0.6)\times10^{-2}$ , which we rescale to our best value $B(D_s^+\rightarrow\phi\pi^+)=(4.5\pm0.4)\times10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(\bar{D}^{*}(2007)^0\rho\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{457}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>0.99±0.11 OUR AVERAGE</b>				
0.97±0.07±0.09		<sup>1,2</sup> DEL-AMO-SA...12	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
1.20 <sup>+0.33</sup> <sub>-0.29</sub> ±0.21		<sup>2</sup> ABE	02w	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.				• • •
1.01±0.10±0.09		<sup>2</sup> AUBERT,B	06s	BABR Repl. by DEL-AMO-SANCHEZ 12
<sup>1</sup> Uses the values of $D$ and $D^*$ branching fractions from PDG 08.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D^{*}(2010)^-\rho\bar{p})/\Gamma_{\text{total}}$				$\Gamma_{458}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>14.5<sup>+3.4</sup><sub>-3.0</sub>±2.7</b>		<sup>1</sup> ANDERSON	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D^-\rho\bar{p}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{459}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>3.32±0.10±0.29</b>		<sup>1,2</sup> DEL-AMO-SA...12	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.				• • •
3.38±0.14±0.29		<sup>2</sup> AUBERT,B	06s	BABR Repl. by DEL-AMO-SANCHEZ 12
<sup>1</sup> Uses the values of $D$ and $D^*$ branching fractions from PDG 08.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(D^{*}(2010)^-\rho\bar{p}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_{460}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>4.7 ± 0.5 OUR AVERAGE</b>				Error includes scale factor of 1.2.
4.55±0.16±0.39		<sup>1,2</sup> DEL-AMO-SA...12	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
6.5 <sup>+1.3</sup> <sub>-1.2</sub> ±1.0		<sup>2</sup> ANDERSON	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc.				• • •
4.81±0.22±0.44		<sup>2</sup> AUBERT,B	06s	BABR Repl. by DEL-AMO-SANCHEZ 12
<sup>1</sup> Uses the values of $D$ and $D^*$ branching fractions from PDG 08.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{D}^0\rho\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{461}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>2.99±0.21±0.45</b>		<sup>1,2</sup> DEL-AMO-SA...12	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses the values of $D$ and $D^*$ branching fractions from PDG 08.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\bar{D}^{*0}\rho\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{462}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT
<b>1.91±0.36±0.29</b>		<sup>1,2</sup> DEL-AMO-SA...12	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses the values of $D$ and $D^*$ branching fractions from PDG 08.				
<sup>2</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\Theta_c\bar{p}\pi^+, \Theta_c\rightarrow D^-\rho)/\Gamma_{\text{total}}$				$\Gamma_{463}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;9</b>	90	<sup>1</sup> AUBERT,B	06s	BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

$\Gamma(\Theta_c^- \bar{p} \pi^+, \Theta_c^- \rightarrow D^{*-} p)/\Gamma_{\text{total}}$					$\Gamma_{464}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;14</b>	90	<sup>1</sup> AUBERT,B	06S	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

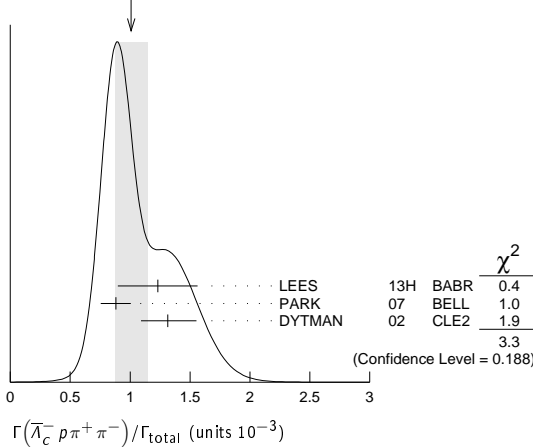
$\Gamma(\bar{\Sigma}_c^{--} \Delta^{++})/\Gamma_{\text{total}}$					$\Gamma_{465}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;8 <math>\times 10^{-4}</math></b>	90	<sup>1</sup> PROCARIO	94	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> PROCARIO 94 reports  $< 0.0012$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Sigma}_c^{--} \Delta^{++})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.043$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 6.35 \times 10^{-2}$ .

$\Gamma(\bar{\Lambda}_c^- p \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{466}/\Gamma$
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
<b>1.01 <math>\pm</math> 0.14 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.			
1.23 $\pm$ 0.05 $\pm$ 0.33		<sup>1,2</sup> LEES	13H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
0.88 $\pm$ 0.11 $\pm$ 0.05		<sup>1,3</sup> PARK	07	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
1.31 $\pm$ 0.21 $\pm$ 0.07		<sup>4</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.87 $\pm$ 0.16 $\pm$ 0.05		<sup>5</sup> GABYSHEV	02	BELL Repl. by PARK 07	
1.33 $\pm$ 0.46 $\pm$ 0.37		<sup>6</sup> FU	97	CLE2 Repl. by DYTMAN 02	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $\Lambda_c^+ \rightarrow p K^- \pi^+$  mode. The second error includes the uncertainty of the branching fraction of the  $\Lambda_c$  decay,  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3)\%$ .  
<sup>3</sup> PARK 07 reports  $(11.2 \pm 0.5 \pm 3.2) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> DYTMAN 02 reports  $(1.67^{+0.27}_{-0.25}) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.05$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>5</sup> GABYSHEV 02 reports  $(1.1 \pm 0.2) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.05$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>6</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching fraction.

WEIGHTED AVERAGE  
1.01 $\pm$ 0.14 (Error scaled by 1.3)



$\Gamma(\bar{\Lambda}_c^- p)/\Gamma_{\text{total}}$					$\Gamma_{467}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>1.52 <math>\pm</math> 0.18 OUR AVERAGE</b>		Error includes scale factor of 1.3.			
1.49 $\pm$ 0.16 $\pm$ 0.08		<sup>1,2</sup> AUBERT	08BN	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
2.19 $\pm$ 0.56 $\pm$ 0.65		<sup>1,3</sup> GABYSHEV	03	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.10 $\pm$ 0.67 $\pm$ 0.77		<sup>1,4</sup> AUBERT	07AV	BABR Repl. by AUBERT 08BN	
< 9	90	<sup>1,5</sup> DYTMAN	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	
< 3.1	90	<sup>1,4</sup> GABYSHEV	02	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	
< 21	90	<sup>6</sup> FU	97	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> AUBERT 08BN reports  $(1.89 \pm 0.21 \pm 0.49) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> The second error for GABYSHEV 03 includes the systematic and the error of  $\Lambda_c \rightarrow \bar{p} K^+ \pi^-$  decay branching fraction.  
<sup>4</sup> Uses the value for  $\Lambda_c \rightarrow p K^- \pi^+$  branching ratio  $(5.0 \pm 1.3)\%$ .  
<sup>5</sup> DYTMAN 02 measurement uses  $\text{B}(\Lambda_c^+ \rightarrow \bar{p} K^+ \pi^-) = 5.0 \pm 1.3\%$ . The second error includes the systematic and the uncertainty of the branching ratio.  
<sup>6</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

$\Gamma(\bar{\Lambda}_c^- p \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{468}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>1.53 <math>\pm</math> 0.17 <math>\pm</math> 0.08</b>		<sup>1,2</sup> AUBERT	10H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<5.9	90	<sup>3</sup> FU	97	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> AUBERT 10H reports  $(1.94 \pm 0.17 \pm 0.52) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p \pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>3</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

$\Gamma(\bar{\Lambda}_c^- p K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{481}/\Gamma$
VALUE (units $10^{-5}$ )		DOCUMENT ID	TECN	COMMENT	
<b>1.97 <math>\pm</math> 0.35 <math>\pm</math> 0.11</b>		<sup>1,2</sup> LEES	15B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> LEES 15B reports  $[(B^0 \rightarrow \bar{\Lambda}_c^- p K^+ K^-)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)] = (12.5 \pm 2.0 \pm 1.0) \times 10^{-7}$  which we divide by our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{\Lambda}_c^- p \phi)/\Gamma_{\text{total}}$					$\Gamma_{482}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;9 <math>\times 10^{-6}</math></b>	90	<sup>1,2</sup> LEES	15B	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> LEES 15B reports  $< 1.2 \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p \phi)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 6.35 \times 10^{-2}$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\Sigma_c(2455)^- p)/\Gamma_{\text{total}}$					$\Gamma_{469}/\Gamma$
VALUE (units $10^{-6}$ )		DOCUMENT ID	TECN	COMMENT	
<b>&lt;24</b>		<sup>1,2</sup> AUBERT	10H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> AUBERT 10H reports  $[\Gamma(B^0 \rightarrow \Sigma_c(2455)^- p)/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)] < 1.5 \times 10^{-6}$  which we divide by our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 6.35 \times 10^{-2}$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\bar{\Lambda}_c^- p \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{470}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;5.07 <math>\times 10^{-3}</math></b>	90	<sup>1</sup> FU	97	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

$\Gamma(\bar{\Lambda}_c^- p \pi^+ \pi^- \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{471}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.74 <math>\times 10^{-3}</math></b>	90	<sup>1</sup> FU	97	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> FU 97 uses PDG 96 values of  $\Lambda_c$  branching ratio.

$\Gamma(\bar{\Lambda}_c^- p \pi^+ \pi^- (\text{nonresonant}))/\Gamma_{\text{total}}$					$\Gamma_{472}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT	
<b>5.4 <math>\pm</math> 1.0 OUR AVERAGE</b>		Error includes scale factor of 1.3.			
7.9 $\pm$ 0.4 $\pm$ 2.0		<sup>1,2</sup> LEES	13H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$	
5.0 $\pm$ 0.8 $\pm$ 0.3		<sup>1,3</sup> PARK	07	BELL $e^+ e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $\Lambda_c^+ \rightarrow p K^- \pi^+$  mode. The second error includes the uncertainty of the branching fraction of the  $\Lambda_c$  decay,  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3)\%$ .  
<sup>3</sup> PARK 07 reports  $(6.4 \pm 0.4 \pm 1.9) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \bar{\Lambda}_c^- p \pi^+ \pi^- (\text{nonresonant}))/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

$B^0$

## $\Gamma(\overline{S}_c(2520)^--\rho\pi^+)/\Gamma_{\text{total}} \quad \Gamma_{473}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.01±0.18 OUR AVERAGE</b>			
1.15±0.10±0.30	1,2 LEES	13H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
0.94±0.21±0.05	1,3 PARK	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.3 ± 0.5 ± 0.1	4 GABYSHEV	02	BELL Repl. by PARK 07

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $\Lambda_c^+ \rightarrow \rho K^- \pi^+$  mode. The second error includes the uncertainty of the branching fraction of the  $\Lambda_c$  decay,  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3)\%$ .  
<sup>3</sup> PARK 07 reports  $(1.2 \pm 0.1 \pm 0.4) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2520)^--\rho\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> GABYSHEV 02 reports  $(1.63^{+0.64}_{-0.58}) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2520)^--\rho\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{S}_c(2520)^0\rho\pi^-)/\Gamma_{\text{total}} \quad \Gamma_{474}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.31 × 10<sup>-4</sup></b>	90	1,2 LEES	13H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.38 × 10 <sup>-4</sup>	90	1 PARK	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<1.21 × 10 <sup>-4</sup>	90	1,2 GABYSHEV	02	BELL Repl. by PARK 07

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses the value for  $\Lambda_c \rightarrow \rho K^- \pi^+$  branching ratio  $(5.0 \pm 1.3)\%$ .

## $\Gamma(\overline{S}_c(2455)^0N^0, N^0 \rightarrow \rho\pi^-)/\Gamma_{\text{total}} \quad \Gamma_{476}/\Gamma$

$N^0$  is the  $N(1440) P_{11}$  or  $N(1535) S_{11}$  or an admixture of the two baryonic states.

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.63±0.16±0.03</b>	1,2 KIM	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> KIM 08 reports  $(0.80 \pm 0.15 \pm 0.25) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^0N^0, N^0 \rightarrow \rho\pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{S}_c(2455)^0\rho\pi^-)/\Gamma_{\text{total}} \quad \Gamma_{475}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.07±0.16 OUR AVERAGE</b>				
0.91±0.07±0.24		1,2 LEES	13H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.10±0.20±0.06		1,3 PARK	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
1.7 ± 0.6 ± 0.1		4 DYTMAN	02	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.38 <sup>+0.36</sup> <sub>-0.32</sub> ±0.02	90	5 GABYSHEV	02	BELL Repl. by PARK 07

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $\Lambda_c^+ \rightarrow \rho K^- \pi^+$  mode. The second error includes the uncertainty of the branching fraction of the  $\Lambda_c$  decay,  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3)\%$ .  
<sup>3</sup> PARK 07 reports  $(1.4 \pm 0.2 \pm 0.4) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^0\rho\pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> DYTMAN 02 reports  $(2.2 \pm 0.7) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^0\rho\pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>5</sup> GABYSHEV 02 reports  $(0.48^{+0.46}_{-0.41}) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^0\rho\pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{S}_c(2455)^--\rho\pi^+)/\Gamma_{\text{total}} \quad \Gamma_{477}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.81±0.24 OUR AVERAGE</b>			
2.13±0.10±0.56	1,2 LEES	13H	BABR $e^+e^- \rightarrow \Upsilon(4S)$
1.65±0.25 <sup>+0.09</sup> <sub>-0.08</sub>	1,3 PARK	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
2.9 ± 0.9 ± 0.2 <sup>+0.2</sup> <sub>-0.1</sub>	4 DYTMAN	02	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.9 <sup>+0.6</sup> <sub>-0.5</sub> ± 0.1	5 GABYSHEV	02	BELL Repl. by PARK 07

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Uses  $\Lambda_c^+ \rightarrow \rho K^- \pi^+$  mode. The second error includes the uncertainty of the branching fraction of the  $\Lambda_c$  decay,  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3)\%$ .  
<sup>3</sup> PARK 07 reports  $(2.1 \pm 0.2 \pm 0.6) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^--\rho\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> DYTMAN 02 reports  $(3.7 \pm 1.1) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^--\rho\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>5</sup> GABYSHEV 02 reports  $(2.38^{+0.75}_{-0.69}) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^--\rho\pi^+)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = 0.05$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## $\Gamma(\overline{\Lambda}_c^-\rho K^+\pi^-)/\Gamma_{\text{total}} \quad \Gamma_{478}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.4±0.7±0.2</b>	1,2 AUBERT	09AG	BABR $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> AUBERT 09AG reports  $(4.33 \pm 0.82 \pm 0.33 \pm 1.13) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{\Lambda}_c^-\rho K^+\pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\overline{S}_c(2455)^--\rho K^+, \overline{S}_c^-- \rightarrow \overline{\Lambda}_c^-\pi^-)/\Gamma_{\text{total}} \quad \Gamma_{479}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.87±0.25<sup>+0.05</sup><sub>-0.04</sub></b>	1,2 AUBERT	09AG	BABR $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> AUBERT 09AG reports  $(1.11 \pm 0.30 \pm 0.09 \pm 0.29) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{S}_c(2455)^--\rho K^+, \overline{S}_c^-- \rightarrow \overline{\Lambda}_c^-\pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\overline{\Lambda}_c^-\rho K^*(892)^0)/\Gamma_{\text{total}} \quad \Gamma_{480}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.42</b>	90	1 AUBERT	09AG	BABR $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\overline{\Lambda}_c^-\rho\overline{p}\rho)/\Gamma_{\text{total}} \quad \Gamma_{483}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.8</b>	1 LEES	14c	BABR $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$  and  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = 0.050 \pm 0.013$ .

## $\Gamma(\overline{\Lambda}_c^-\Lambda K^+)/\Gamma_{\text{total}} \quad \Gamma_{484}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.8±1.1<sup>+0.2</sup><sub>-0.3</sub></b>	1,2 LEES	11F	BABR $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  from Upsilon(4S) decays.  
<sup>2</sup> LEES 11F reports  $(3.8 \pm 0.8 \pm 0.2 \pm 1.0) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \overline{\Lambda}_c^-\Lambda K^+)/\Gamma_{\text{total}}] / [B(\Lambda_c^+ \rightarrow \rho K^- \pi^+)] / [B(\Lambda \rightarrow p\pi^-)]$  assuming  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ ,  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5) \times 10^{-2}$ , which we rescale to our best values  $B(\Lambda_c^+ \rightarrow \rho K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ ,  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values. The reported uncertainties are statistical, systematic, and  $\overline{\Lambda}_c^-$  branching fraction uncertainty.

## $\Gamma(\overline{\Lambda}_c^-\Lambda_c^+)/\Gamma_{\text{total}} \quad \Gamma_{485}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.6</b>	95	1 AAIJ	14AA	LHCb $pp$ at 7 TeV

- • • We do not use the following data for averages, fits, limits, etc. • • •  
<6.2 90 2 UCHIDA 08 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Uses  $B(\overline{B}^0 \rightarrow D^+D_s^-) = (7.2 \pm 0.8) \times 10^{-3}$ .  
<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

## $\Gamma(\overline{\Lambda}_c(2593)^- / \overline{\Lambda}_c(2625)^-\rho)/\Gamma_{\text{total}} \quad \Gamma_{486}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 × 10<sup>-4</sup></b>	90	1,2 DYTMAN	02	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> DYTMAN 02 measurement uses  $B(\Lambda_c^- \rightarrow \overline{p}K^+\pi^-) = 5.0 \pm 1.3\%$ . The second error includes the systematic and the uncertainty of the branching ratio.



See key on page 601

Meson Particle Listings  
 $B^0$  $\Gamma(\Xi_c^- \Lambda_c^+ \cdot \Xi_c^- \rightarrow \Xi^+ \pi^- \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{487}/\Gamma$ 

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.7 \pm 1.8</math> OUR AVERAGE</b>	Error includes scale factor of 2.2.		
$1.2 \pm 0.9 \pm 0.1$	1,2 AUBERT	08H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$7.3 \pm 3.3 \pm 0.4$	2,3 CHISTOV	06A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 08H reports  $(1.5 \pm 1.07 \pm 0.44) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \Xi_c^- \Lambda_c^+ \cdot \Xi_c^- \rightarrow \Xi^+ \pi^- \pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> CHISTOV 06A reports  $(9.3 \pm 3.7 \pm 3.1) \times 10^{-5}$  from a measurement of  $[\Gamma(B^0 \rightarrow \Xi_c^- \Lambda_c^+ \cdot \Xi_c^- \rightarrow \Xi^+ \pi^- \pi^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\Lambda_c^+ \Lambda_c^- K^0)/\Gamma_{\text{total}}$   $\Gamma_{488}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>4.3 \pm 2.2</math> OUR AVERAGE</b>			
$3.0 \pm 2.8 \pm 0.2$	1,2 AUBERT	08H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$6.2 \pm 3.7 \pm 0.3$	2,3 GABYSHEV	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 08H reports  $(0.38 \pm 0.31 \pm 0.21) \times 10^{-3}$  from a measurement of  $[\Gamma(B^0 \rightarrow \Lambda_c^+ \Lambda_c^- K^0)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> GABYSHEV 06 reports  $(7.9 \pm 2.9 \pm 4.3) \times 10^{-4}$  from a measurement of  $[\Gamma(B^0 \rightarrow \Lambda_c^+ \Lambda_c^- K^0)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$   $\Gamma_{489}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.2 \times 10^{-7}</math></b>	90	1 DEL-AMO-SA...11A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 6.2 \times 10^{-7}$	90	1 VILLA	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.7 \times 10^{-6}$	90	1 AUBERT	01i BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 3.9 \times 10^{-5}$	90	2 ACCIARRI	95i L3	$e^+ e^- \rightarrow Z$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ACCIARRI 95i assumes  $f_{B^0} = 39.5 \pm 4.0$  and  $f_{B_s} = 12.0 \pm 3.0\%$ .

 $\Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{490}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 8.3 \times 10^{-8}</math></b>	90	AALTONEN	09P CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 11.3 \times 10^{-8}$	90	1 AUBERT	08P BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 6.1 \times 10^{-8}$	90	1 AUBERT	05W BABR	Repl. by AUBERT 08P
$< 1.9 \times 10^{-7}$	90	1 CHANG	03 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 8.3 \times 10^{-7}$	90	1 BERGFELD	00B CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.4 \times 10^{-5}$	90	2 ACCIARRI	97B L3	$e^+ e^- \rightarrow Z$
$< 5.9 \times 10^{-6}$	90	AMMAR	94 CLE2	Repl. by BERGFELD 00B
$< 2.6 \times 10^{-5}$	90	3 AVERY	89B CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 7.6 \times 10^{-5}$	90	4 ALBRECHT	87D ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 6.4 \times 10^{-5}$	90	5 AVERY	87 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 3 \times 10^{-4}$	90	GILES	84 CLEO	Repl. by AVERY 87

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> ACCIARRI 97B assume PDG 96 production fractions for  $B^+$ ,  $B^0$ ,  $B_s$ , and  $\Lambda_b$ .

<sup>3</sup> AVERY 89B reports  $< 3 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>4</sup> ALBRECHT 87D reports  $< 8.5 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>5</sup> AVERY 87 reports  $< 8 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

 $\Gamma(e^+ e^- \gamma)/\Gamma_{\text{total}}$   $\Gamma_{491}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.2 \times 10^{-7}</math></b>	90	AUBERT	08c BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $\Gamma(\mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{492}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.39 \pm 0.16</math> <math>-0.14</math></b>		1 KHACHATRY...15BE	LHC	$p\bar{p}$ at 7, 8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 0.80$	90	2 AAIJ	13B LHCB	Repl. by AAIJ 13BA
$< 0.63$	90	3 AAIJ	13BA LHCB	Repl. by KHACHA-TRYAN 15BE
$< 3.8$	90	4 AALTONEN	13F CDF	$p\bar{p}$ at 1.96 TeV
$< 0.92$	90	5 CHATRCHYAN	13AW CMS	$p\bar{p}$ at 7, 8 TeV
$< 2.6$	90	2 AAIJ	12A LHCB	Repl. by AAIJ 12W
$< 0.81$	90	6 AAIJ	12W LHCB	Repl. by AAIJ 13B
$< 1.4$	90	6 CHATRCHYAN	12A CMS	$p\bar{p}$ at 7 TeV
$< 12$	90	7 AAIJ	11B LHCB	Repl. by AAIJ 12A
$< 5.0$	90	6 AALTONEN	11AG CDF	$p\bar{p}$ at 1.96 TeV
$< 3.7$	90	6 CHATRCHYAN	11T CMS	Repl. by CHATRCHYAN 12A
$< 15$	90	8 AALTONEN	08i CDF	Repl. by AALTONEN 11AG
$< 52$	90	9 AUBERT	08P BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 39$	90	10 ABULENCIA	05 CDF	Repl. by AALTONEN 08i
$< 83$	90	9 AUBERT	05W BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 150$	90	11 ACOSTA	04D CDF	$p\bar{p}$ at 1.96 TeV
$< 160$	90	9 CHANG	03 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 610$	90	9 BERGFELD	00B CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 40000$	90	ABBOTT	98B D0	$p\bar{p}$ 1.8 TeV
$< 680$	90	12 ABE	98 CDF	$p\bar{p}$ at 1.8 TeV
$< 10000$	90	13 ACCIARRI	97B L3	$e^+ e^- \rightarrow Z$
$< 1600$	90	14 ABE	96L CDF	Repl. by ABE 98
$< 5900$	90	AMMAR	94 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 8300$	90	15 ALBAJAR	91c UA1	$E_{\text{cm}}^{p\bar{p}} = 630$ GeV
$< 12000$	90	16 ALBAJAR	91c UA1	$E_{\text{cm}}^{p\bar{p}} = 630$ GeV
$< 43000$	90	17 AVERY	89B CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 45000$	90	18 ALBRECHT	87D ARG	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 77000$	90	19 AVERY	87 CLEO	$e^+ e^- \rightarrow \Upsilon(4S)$
$< 200000$	90	GILES	84 CLEO	Repl. by AVERY 87

<sup>1</sup> Derived from the combined fit to CMS and LHCb data. Uncertainty includes both statistical and systematic component. Also reports  $B(B^0 \rightarrow \mu^+ \mu^-)/B(B_S \rightarrow \mu^+ \mu^-) = 0.14 \pm 0.08$ .

<sup>2</sup> Uses  $B(B^+ \rightarrow J/\psi K^+ \rightarrow \mu^+ \mu^- K^+) = (6.01 \pm 0.21) \times 10^{-5}$  and  $B(B^0 \rightarrow K^+ \pi^-) = (1.94 \pm 0.06) \times 10^{-5}$  for normalization.

<sup>3</sup> Reports also a limit of  $< 7.4 \times 10^{-10}$  at 95% CL. Uses normalization modes  $B^+ \rightarrow J/\psi K^+ \rightarrow \mu^+ \mu^- K^+$  and  $B^0 \rightarrow K^+ \pi^-$ .

<sup>4</sup> Uses normalization mode  $B(B^+ \rightarrow J/\psi K^+) = (10.22 \pm 0.35) \times 10^{-4}$ .

<sup>5</sup> Uses  $B(B^+ \rightarrow J/\psi K^+ \rightarrow \mu^+ \mu^- K^+) = (6.0 \pm 0.2) \times 10^{-5}$  for normalization.

<sup>6</sup> Uses  $B(B^+ \rightarrow J/\psi K^+ \rightarrow \mu^+ \mu^- K^+) = (6.01 \pm 0.21) \times 10^{-5}$ .

<sup>7</sup> Uses  $B$  production ratio  $f(\bar{D} \rightarrow B^+)/f(\bar{D} \rightarrow B_S^0) = 3.71 \pm 0.47$  and three normalization modes.

<sup>8</sup> Uses  $B(B^+ \rightarrow J/\psi K^+) B(J/\psi \rightarrow \mu^+ \mu^-) = (5.94 \pm 0.21) \times 10^{-5}$ .

<sup>9</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>10</sup> Uses  $B(B^+ \rightarrow J/\psi K^+) B(J/\psi \rightarrow \mu^+ \mu^-) = (5.88 \pm 0.26) \times 10^{-5}$ .

<sup>11</sup> Assumes production cross-section  $\sigma(B_S)/\sigma(B^+) = 0.100/0.391$  and the CDF measured value of  $\sigma(B^+) = 3.6 \pm 0.6 \mu\text{b}$ .

<sup>12</sup> ABE 98 assumes production of  $\sigma(B^0) = \sigma(B^+)$  and  $\sigma(B_S)/\sigma(B^0) = 1/3$ . They normalize to their measured  $\sigma(B^0, p_T(B) > 6, |y| < 1.0) = 2.39 \pm 0.32 \pm 0.44 \mu\text{b}$ .

<sup>13</sup> ACCIARRI 97B assume PDG 96 production fractions for  $B^+$ ,  $B^0$ ,  $B_s$ , and  $\Lambda_b$ .

<sup>14</sup> ABE 96i assumes equal  $B^0$  and  $B^+$  production. They normalize to their measured  $\sigma(B^+, p_T(B) > 6 \text{ GeV}/c, |y| < 1) = 2.39 \pm 0.54 \mu\text{b}$ .

<sup>15</sup>  $B^0$  and  $B_S^0$  are not separated.

<sup>16</sup> Obtained from unseparated  $B^0$  and  $B_S^0$  measurement by assuming a  $B^0:B_S^0$  ratio 2:1.

<sup>17</sup> AVERY 89B reports  $< 5 \times 10^{-3}$  assuming the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>18</sup> ALBRECHT 87D reports  $< 5 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0 \bar{B}^0$ . We rescale to 50%.

<sup>19</sup> AVERY 87 reports  $< 9 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

 $\Gamma(\mu^+ \mu^- \gamma)/\Gamma_{\text{total}}$   $\Gamma_{493}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.6 \times 10^{-7}</math></b>	90	AUBERT	08c BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $\Gamma(\tau^+ \tau^-)/\Gamma_{\text{total}}$   $\Gamma_{496}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.1 \times 10^{-3}</math></b>	90	1 AUBERT	06s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
1 Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

 $\Gamma(\mu^+ \mu^- \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{494}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 5.3 \times 10^{-9}</math></b>	90	1 AAIJ	13AW LHCB	$p\bar{p}$ at 7 TeV

<sup>1</sup> Also reports a limit of  $< 6.6 \times 10^{-9}$  at 95% CL.

 $\Gamma(SP, S \rightarrow \mu^+ \mu^-, P \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{495}/\Gamma$ 

Here  $S$  and  $P$  are the hypothetical scalar and pseudoscalar particles with masses of 2.5 GeV/ $c^2$  and 214.3 MeV/ $c^2$ , respectively.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 5.1 \times 10^{-9}</math></b>	90	1 AAIJ	13AW LHCB	$p\bar{p}$ at 7 TeV

<sup>1</sup> Also reports a limit of  $< 6.3 \times 10^{-9}$  at 95% CL.

# Meson Particle Listings

$B^0$

$\Gamma(\pi^0 \ell^+ \ell^-)/\Gamma_{\text{total}}$					$\Gamma_{497}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;5.3 \times 10^{-8}</math></b>	90	<sup>1</sup> LEES	13M	BABR $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$<1.5 \times 10^{-7}$	90	<sup>1</sup> WEI	08A	BELL $e^+e^- \rightarrow \gamma(4S)$	
$<1.2 \times 10^{-7}$	90	<sup>1</sup> AUBERT	07AG	BABR Repl. by LEES 13M	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(\pi^0 \nu \bar{\nu})/\Gamma_{\text{total}}$					$\Gamma_{503}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interaction.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6.9 \times 10^{-5}</math></b>	90	<sup>1</sup> LUTZ	13	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$<2.2 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D	BELL Repl. by LUTZ 13	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{498}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;8.4 \times 10^{-8}</math></b>	90	<sup>1</sup> LEES	13M	BABR $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$<2.3 \times 10^{-7}$	90	<sup>1</sup> WEI	08A	BELL $e^+e^- \rightarrow \gamma(4S)$	
$<1.4 \times 10^{-7}$	90	<sup>1</sup> AUBERT	07AG	BABR Repl. by LEES 13M	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(\pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{499}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6.9 \times 10^{-8}</math></b>	90	<sup>1</sup> LEES	13M	BABR $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$<1.8 \times 10^{-7}$	90	<sup>1</sup> WEI	08A	BELL $e^+e^- \rightarrow \gamma(4S)$	
$<5.1 \times 10^{-7}$	90	<sup>1</sup> AUBERT	07AG	BABR $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(\eta \ell^+ \ell^-)/\Gamma_{\text{total}}$					$\Gamma_{500}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2.4 \times 10^{-8}</math></b>	90	<sup>1</sup> LEES	13M	BABR $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{501}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;10.8 \times 10^{-8}</math></b>	90	<sup>1</sup> LEES	13M	BABR $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(\eta \mu^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{502}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;11.2 \times 10^{-8}</math></b>	90	<sup>1</sup> LEES	13M	BABR $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(K^0 \ell^+ \ell^-)/\Gamma_{\text{total}}$					$\Gamma_{504}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.1^{+0.8}_{-0.7}</math> OUR AVERAGE</b>					
$2.1^{+1.5}_{-1.3} \pm 0.2$		<sup>1</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$	
$3.4^{+0.9}_{-0.8} \pm 0.2$		<sup>1</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$2.9^{+1.6}_{-1.3} \pm 0.3$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T	
$<6.8$	90	<sup>1</sup> ISHIKAWA	03	BELL $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\gamma(4S)$ .

$\Gamma(K^0 e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{505}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interactions.					
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.6^{+1.0}_{-0.8}</math> OUR AVERAGE</b>					
$0.8^{+1.5}_{-1.2} \pm 0.1$		<sup>1</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$	
$2.0^{+1.4}_{-1.0} \pm 0.1$		<sup>1</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$1.3^{+1.6}_{-1.1} \pm 0.2$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T	
— $2.1^{+2.3}_{-1.6} \pm 0.8$		<sup>1</sup> AUBERT	03U	BABR $e^+e^- \rightarrow \gamma(4S)$	
$< 5.4$	90	<sup>2</sup> ISHIKAWA	03	BELL $e^+e^- \rightarrow \gamma(4S)$	
$< 27$	90	<sup>1</sup> ABE	02	BELL Repl. by ISHIKAWA 03	
$< 38$	90	<sup>1</sup> AUBERT	02L	BABR $e^+e^- \rightarrow \gamma(4S)$	
$< 84.5$	90	<sup>3</sup> ANDERSON	01B	CLE2 $e^+e^- \rightarrow \gamma(4S)$	
$< 3000$	90	<sup>1</sup> ALBRECHT	91E	ARG $e^+e^- \rightarrow \gamma(4S)$	
$< 5200$	90	<sup>4</sup> AVERY	87	CLEO $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\gamma(4S)$ .

<sup>3</sup> The result is for di-lepton masses above 0.5 GeV.

<sup>4</sup> AVERY 87 reports  $< 6.5 \times 10^{-4}$  assuming the  $\gamma(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

$\Gamma(K^0 \nu \bar{\nu})/\Gamma_{\text{total}}$					$\Gamma_{507}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interaction.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 4.9 \times 10^{-5}</math></b>	90	<sup>1,2</sup> LEES	13i	BABR $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$<19.4 \times 10^{-5}$	90	<sup>1</sup> LUTZ	13	BELL $e^+e^- \rightarrow \gamma(4S)$	
$< 5.6 \times 10^{-5}$	90	<sup>1</sup> DEL-AMO-SA...	10Q	BABR Repl. by LEES 13i	
$< 1.6 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D	BELL $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

<sup>2</sup> Also reported a limit  $< 8.1 \times 10^{-5}$  at 90% CL obtained using a fully reconstructed hadronic  $B$ -tag evnets.

$\Gamma(\rho^0 \nu \bar{\nu})/\Gamma_{\text{total}}$					$\Gamma_{508}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interaction.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2.08 \times 10^{-4}</math></b>	90	<sup>1</sup> LUTZ	13	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$<4.4 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D	BELL Repl. by LUTZ 13	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

$\Gamma(K^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{506}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interactions.					
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>3.39 \pm 0.34</math> OUR FIT</b>					
<b><math>3.4 \pm 0.4</math> OUR AVERAGE</b>					
$3.27 \pm 0.34 \pm 0.17$		<sup>1</sup> AAIJ	14M	LHCb $p\bar{p}$ at 7, 8 TeV	
$4.9^{+2.9}_{-2.5} \pm 0.3$		<sup>2</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$	
$4.4^{+1.3}_{-1.1} \pm 0.3$		<sup>2</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$3.1^{+0.7}_{-0.6}$		AAIJ	12AH	LHCb Repl. by AAIJ 14M	
$5.9^{+3.3}_{-2.6} \pm 0.7$		<sup>2</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T	
$1.63^{+0.82}_{-0.63} \pm 0.14$		<sup>2</sup> AUBERT	03U	BABR Repl. by AUBERT,B 06J	
$5.6^{+2.9}_{-2.3} \pm 0.5$		<sup>3</sup> ISHIKAWA	03	BELL Repl. by WEI 09A	
$<33$	90	<sup>2</sup> ABE	02	BELL Repl. by ISHIKAWA 03	
$<36$	90	AUBERT	02L	BABR $e^+e^- \rightarrow \gamma(4S)$	
$<66.4$	90	<sup>4</sup> ANDERSON	01B	CLE2 $e^+e^- \rightarrow \gamma(4S)$	
$<5200$	90	ALBRECHT	91E	ARG $e^+e^- \rightarrow \gamma(4S)$	
$<3600$	90	<sup>5</sup> AVERY	87	CLEO $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Uses  $B(B^0 \rightarrow J/\psi(1S) K^0) = (0.928 \pm 0.013 \pm 0.037) \times 10^{-3}$  for normalization.

<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\gamma(4S)$ .

<sup>3</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\gamma(4S)$ . The second error is a total of systematic uncertainties including model dependence.

<sup>4</sup> The result is for di-lepton masses above 0.5 GeV.

<sup>5</sup> AVERY 87 reports  $< 4.5 \times 10^{-4}$  assuming the  $\gamma(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

$\Gamma(K^0 \mu^+ \mu^-)/\Gamma(J/\psi(1S) K^0)$					$\Gamma_{506}/\Gamma_{183}$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.39 \pm 0.04</math> OUR FIT</b>					
<b><math>0.37 \pm 0.12 \pm 0.02</math></b>		AALTONEN	11Ai	CDF $p\bar{p}$ at 1.96 TeV	

$\Gamma(K^*(892)^0 \ell^+ \ell^-)/\Gamma_{\text{total}}$					$\Gamma_{509}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interactions.					
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>9.9^{+1.2}_{-1.1}</math> OUR AVERAGE</b>					
$10.3^{+2.2}_{-2.1} \pm 0.7$		<sup>1</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$	
$9.7^{+1.3}_{-1.1} \pm 0.7$		<sup>1</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$8.1^{+2.1}_{-1.9} \pm 0.9$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T	
$11.7^{+3.0}_{-2.7} \pm 0.9$		<sup>1</sup> ISHIKAWA	03	BELL Repl. by WEI 09A	

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\gamma(4S)$ .

$\Gamma(K^*(892)^0 e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_{510}/\Gamma$
Test for $\Delta B=1$ weak neutral current. Allowed by higher-order electroweak interactions.					
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>10.3^{+1.9}_{-1.7}</math> OUR AVERAGE</b>					
$8.6^{+2.6}_{-2.4} \pm 0.5$		<sup>1</sup> AUBERT	09T	BABR $e^+e^- \rightarrow \gamma(4S)$	
$11.8^{+2.7}_{-2.2} \pm 0.9$		<sup>1</sup> WEI	09A	BELL $e^+e^- \rightarrow \gamma(4S)$	
• • • We do not use the following data for averages, fits, limits, etc.				• • •	
$10.4^{+3.3}_{-2.9} \pm 1.1$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T	
$11.1^{+5.6}_{-4.7} \pm 1.1$		<sup>1</sup> AUBERT	03U	BABR $e^+e^- \rightarrow \gamma(4S)$	
$< 24$	90	<sup>2</sup> ISHIKAWA	03	BELL $e^+e^- \rightarrow \gamma(4S)$	
$< 64$	90	<sup>1</sup> ABE	02	BELL Repl. by ISHIKAWA 03	
$< 67$	90	<sup>1</sup> AUBERT	02L	BABR $e^+e^- \rightarrow \gamma(4S)$	
$<2900$	90	ALBRECHT	91E	ARG $e^+e^- \rightarrow \gamma(4S)$	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{511}/\Gamma$   
 Test for  $\Delta B=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>10.2 \pm 0.9</math> OUR FIT</b>				
<b><math>11.1^{+1.8}_{-1.4}</math> OUR AVERAGE</b>				
$13.5^{+4.0}_{-3.7} \pm 1.0$		<sup>1</sup> AUBERT	09T	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$10.6^{+1.9}_{-1.4} \pm 0.7$		<sup>1</sup> WEI	09A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.7^{+3.8}_{-3.3} \pm 1.2$		<sup>1</sup> AUBERT,B	06J	BABR Repl. by AUBERT 09T
$8.6^{+7.9}_{-5.8} \pm 1.1$		<sup>1</sup> AUBERT	03U	BABR Repl. by AUBERT,B 06J
$13.3^{+4.2}_{-3.7} \pm 1.1$		<sup>2</sup> ISHIKAWA	03	BELL Repl. by WEI 09A
< 42	90	<sup>1</sup> ABE	02	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< 33	90	AUBERT	02L	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< 40	90	<sup>3</sup> AFFOLDER	99B	CDF $p\bar{p}$ at 1.8 TeV
< 250	90	<sup>4</sup> ABE	96L	CDF Repl. by AFFOLDER 99B
< 230	90	<sup>5</sup> ALBAJAR	91c	UA1 $E_{\text{cm}}^{\text{pp}} = 630$ GeV
< 3400	90	ALBRECHT	91E	ARG $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ . The second error is a total of systematic uncertainties including model dependence.<sup>3</sup> AFFOLDER 99B measured relative to  $B^0 \rightarrow J/\psi(1S) K^*(892)^0$ .<sup>4</sup> ABE 96L measured relative to  $B^0 \rightarrow J/\psi(1S) K^*(892)^0$  using PDG 94 branching ratios.<sup>5</sup> ALBAJAR 91C assumes 36% of  $\bar{D}$  quarks give  $B^0$  mesons.

$\Gamma(K^*(892)^0 \mu^+ \mu^-)/\Gamma(J/\psi(1S) K^*(892)^0)$   $\Gamma_{511}/\Gamma_{185}$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.80 \pm 0.07</math> OUR FIT</b>			
<b><math>0.77 \pm 0.08 \pm 0.03</math></b>	AALTONEN	11A	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.80 \pm 0.10 \pm 0.06$	AALTONEN	11L	CDF Repl. by AALTONEN 11A
$0.61 \pm 0.23 \pm 0.07$	AALTONEN	09B	CDF Repl. by AALTONEN 11L

$\Gamma(K^*(892)^0 \chi, \chi \rightarrow \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{512}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< $\sim 10^{-9}$	95	<sup>1</sup> AAJ	15A	LHCB $pp$ at 7, 8 TeV

<sup>1</sup> The limit is obtained as a function of di-muon mass. A normalizing mode branching fraction value of  $B(B^0 \rightarrow K^{*0} \mu^+ \mu^-) = (1.6 \pm 0.3) \times 10^{-7}$  is used.

$\Gamma(\pi^+ \pi^- \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{513}/\Gamma$

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.1 \pm 0.5 \pm 0.1</math></b>	<sup>1</sup> AAJ	15s	LHCB $pp$ at 7, 8 TeV

<sup>1</sup> AAJ 15s reports  $(2.11 \pm 0.51 \pm 0.15 \pm 0.16) \times 10^{-8}$  from a measurement of  $[\Gamma(B^0 \rightarrow \pi^+ \pi^- \mu^+ \mu^-)/\Gamma_{\text{total}}] / [B(B^0 \rightarrow J/\psi(1S) K^*(892)^0)]$  assuming  $B(B^0 \rightarrow J/\psi(1S) K^*(892)^0) = (1.3 \pm 0.1) \times 10^{-3}$ , which we rescale to our best value  $B(B^0 \rightarrow J/\psi(1S) K^*(892)^0) = (1.28 \pm 0.05) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^*(892)^0 \nu \bar{\nu})/\Gamma_{\text{total}}$   $\Gamma_{514}/\Gamma$

Test for  $\Delta B=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 5.5 \times 10^{-5}</math></b>	90	<sup>1</sup> LUTZ	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.2 \times 10^{-4}$	90	<sup>1,2</sup> LEES	13i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.2 \times 10^{-4}$	90	AUBERT	08B	BABR Repl. by LEES 13i
$< 3.4 \times 10^{-4}$	90	<sup>1</sup> CHEN	07D	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$< 1.0 \times 10^{-3}$	90	<sup>3</sup> ADAM	96D	DLPH $e^+ e^- \rightarrow Z$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>2</sup> Also reported a limit  $< 9.3 \times 10^{-5}$  at 90% CL obtained using a fully reconstructed hadronic  $B$ -tag evnets.<sup>3</sup> ADAM 96D assumes  $f_{B^0} = f_{B^-} = 0.39$  and  $f_{B_s} = 0.12$ .

$\Gamma(\phi \nu \bar{\nu})/\Gamma_{\text{total}}$   $\Gamma_{515}/\Gamma$

Test for  $\Delta B=1$  weak neutral current. Allowed by higher-order electroweak interaction.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.27 \times 10^{-4}</math></b>	90	<sup>1</sup> LUTZ	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 5.8 \times 10^{-5}$	90	<sup>1</sup> CHEN	07D	BELL Repl. by LUTZ 13

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{516}/\Gamma$

Test of lepton family number conservation. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.8 \times 10^{-9}</math></b>	90	<sup>1</sup> AAJ	13BMLHCB	$pp$ at 7 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< $6.4 \times 10^{-8}$	90	AALTONEN	09P	CDF $p\bar{p}$ at 1.96 TeV
< $9.2 \times 10^{-8}$	90	<sup>2</sup> AUBERT	08P	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< $1.8 \times 10^{-7}$	90	<sup>2</sup> AUBERT	05W	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
< $1.7 \times 10^{-7}$	90	<sup>2</sup> CHANG	03	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< $15 \times 10^{-7}$	90	<sup>2</sup> BERGFELD	00B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
< $3.5 \times 10^{-6}$	90	ABE	98V	CDF $p\bar{p}$ at 1.8 TeV
< $1.6 \times 10^{-5}$	90	<sup>3</sup> ACCIARRI	97B	L3 $e^+ e^- \rightarrow Z$
< $5.9 \times 10^{-6}$	90	AMMAR	94	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
< $3.4 \times 10^{-5}$	90	<sup>4</sup> AVERY	89B	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
< $4.5 \times 10^{-5}$	90	<sup>5</sup> ALBRECHT	87D	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
< $7.7 \times 10^{-5}$	90	<sup>6</sup> AVERY	87	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
< $3 \times 10^{-4}$	90	GILES	84	CLEO Repl. by AVERY 87

<sup>1</sup> Uses normalization mode  $B(B^0 \rightarrow K^+ \pi^-) = (19.4 \pm 0.6) \times 10^{-6}$ .<sup>2</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .<sup>3</sup> ACCIARRI 97B assume PDG 96 production fractions for  $B^+$ ,  $B^0$ ,  $B_s$ , and  $\Lambda_b$ .<sup>4</sup> Paper assumes the  $\Upsilon(4S)$  decays 43% to  $B^0 \bar{B}^0$ . We rescale to 50%.<sup>5</sup> ALBRECHT 87D reports  $< 5 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 45% to  $B^0 \bar{B}^0$ . We rescale to 50%.<sup>6</sup> AVERY 87 reports  $< 9 \times 10^{-5}$  assuming the  $\Upsilon(4S)$  decays 40% to  $B^0 \bar{B}^0$ . We rescale to 50%.

$\Gamma(\pi^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{517}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.4 \times 10^{-7}</math></b>	90	<sup>1</sup> AUBERT	07AG	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{518}/\Gamma$

Test of lepton family number conservation.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.7</b>	90	<sup>1</sup> AUBERT,B	06i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 40	90	<sup>1</sup> AUBERT	02L	BABR Repl. by AUBERT,B 06J
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0 e^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{519}/\Gamma$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 5.3</b>	90	<sup>1</sup> AUBERT,B	06i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0 e^- \mu^+)/\Gamma_{\text{total}}$   $\Gamma_{520}/\Gamma$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.4</b>	90	<sup>1</sup> AUBERT,B	06i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ .

$\Gamma(K^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{521}/\Gamma$

Test of lepton family number conservation.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 5.8</b>	90	<sup>1</sup> AUBERT,B	06i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 34	90	<sup>1</sup> AUBERT	02L	BABR Repl. by AUBERT,B 06J
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<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(e^\pm \tau^\mp)/\Gamma_{\text{total}}$   $\Gamma_{522}/\Gamma$

Test of lepton family number conservation. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.8 \times 10^{-5}</math></b>	90	<sup>1</sup> AUBERT	08AD	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< $1.1 \times 10^{-4}$	90	BORNHEIM	04	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
< $5.3 \times 10^{-4}$	90	AMMAR	94	CLE2 Repl. by BORNHEIM 04

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\mu^\pm \tau^\mp)/\Gamma_{\text{total}}$   $\Gamma_{523}/\Gamma$

Test of lepton family number conservation. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.2 \times 10^{-5}</math></b>	90	<sup>1</sup> AUBERT	08AD	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< $3.8 \times 10^{-5}$	90	BORNHEIM	04	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
< $8.3 \times 10^{-4}$	90	AMMAR	94	CLE2 Repl. by BORNHEIM 04

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\text{invisible})/\Gamma_{\text{total}}$   $\Gamma_{524}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.4</b>	90	<sup>1</sup> LEES	12T	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 13	90	<sup>2</sup> HSU	12	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
< 22	90	<sup>1</sup> AUBERT,B	04i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the fully reconstructed  $B^0 \rightarrow D^{(*)} - \ell^+ \nu_\ell$  events as a tag.<sup>2</sup> Identified by fully reconstructing a hadronic decay of the accompanying  $B$  meson and requiring no other particles in the event.

## Meson Particle Listings

 $B^0$  $\Gamma(\nu\bar{\nu}\gamma)/\Gamma_{\text{total}}$   $\Gamma_{525}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.7	90	<sup>1</sup> LEES	12T	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<4.7	90	<sup>1</sup> AUBERT,B	04J	BABR Repl. by LEES 12T
<sup>1</sup> Uses the fully reconstructed $B^0 \rightarrow D(*)-\ell^+\nu_\ell$ events as a tag.				

 $\Gamma(\Lambda_c^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{526}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<1.4 $\times 10^{-6}$	90	<sup>1,2</sup> DEL-AMO-SA..11K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> DEL-AMO-SANCHEZ 11K reports $< 180 \times 10^{-8}$ from a measurement of $[\Gamma(B^0 \rightarrow \Lambda_c^+\mu^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$ assuming $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 6.35 \times 10^{-2}$ .				
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^+B^-) = (48.4 \pm 0.6)\%$ .				

 $\Gamma(\Lambda_c^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{527}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<4 $\times 10^{-6}$	90	<sup>1,2</sup> DEL-AMO-SA..11K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> DEL-AMO-SANCHEZ 11K reports $< 520 \times 10^{-8}$ from a measurement of $[\Gamma(B^0 \rightarrow \Lambda_c^+e^-)/\Gamma_{\text{total}}] \times [B(\Lambda_c^+ \rightarrow pK^-\pi^+)]$ assuming $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value $B(\Lambda_c^+ \rightarrow pK^-\pi^+) = 6.35 \times 10^{-2}$ .				
<sup>2</sup> Uses $B(\Upsilon(4S) \rightarrow B^0\bar{B}^0) = (51.6 \pm 0.6)\%$ and $B(\Upsilon(4S) \rightarrow B^+B^-) = (48.4 \pm 0.6)\%$ .				

 $B_s^0$  CROSS-PARTICLE BRANCHING RATIOS $\Gamma([K^+K^-]_D K^*(892)^0)/\Gamma_{\text{total}} \times B(B_s^0 \rightarrow [K^+K^-]_D K^*(892)^0)$   $\Gamma_{142}/\Gamma \times B$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.10 \pm 0.02 \pm 0.01$	AAIJ	14BN LHCB	$pp$ at 7, 8 TeV

 $\Gamma([\pi^+\pi^-]_D K^*(892)^0)/\Gamma_{\text{total}} \times B(B_s^0 \rightarrow [\pi^+\pi^-]_D K^*(892)^0)$   $\Gamma_{143}/\Gamma \times B$ 

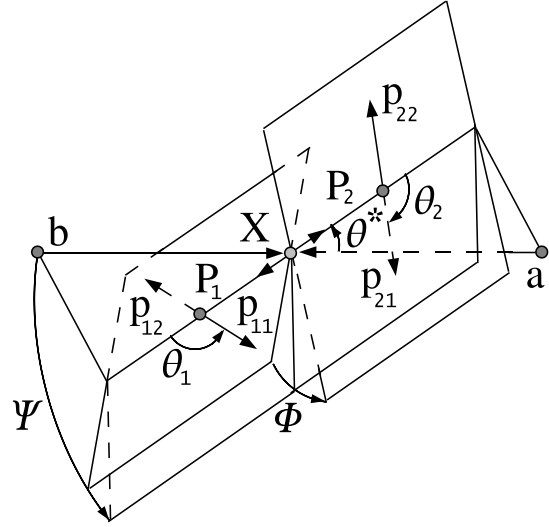
VALUE	DOCUMENT ID	TECN	COMMENT
$0.15 \pm 0.04 \pm 0.01$	AAIJ	14BN LHCB	$pp$ at 7, 8 TeV

POLARIZATION IN  $B$  DECAYS

Revised August 2015 by A. V. Gritsan (Johns Hopkins University).

We review the notation used in polarization measurements in particle production and decay, with a particular emphasis on the  $B$  decays and the  $CP$ -violating observables in polarization measurements. We look at several examples of vector-vector and vector-tensor  $B$  meson decays, while more details about the theory and experimental results in  $B$  decays can be found in a separate mini-review [1] in this *Review*.

Figure 1 illustrates angular observables in an example of the sequential process  $ab \rightarrow X \rightarrow P_1P_2 \rightarrow (p_{11}p_{12})(p_{21}p_{22})$  [2]. The angular distributions are of particular interest because they are sensitive to spin correlations and reveal properties of particles and their interactions, such as quantum numbers and couplings. In the case of a spin-zero particle  $X$ , such as  $B$  meson or a Higgs boson, there are no spin correlations in the production mechanism and the decay chain is to be analyzed. The angular distribution of decay products can be expressed as a function of three helicity angles which describe the alignment of the particles in the decay chain. The analyzer of the  $B$ -daughter polarization is normally chosen for two-body decays, as the direction of the daughters in the center-of-mass of the parent (*e.g.*,  $\rho \rightarrow 2\pi$ ) [3], and for three-body decays as the normal to the decay plane (*e.g.*,  $\omega \rightarrow 3\pi$ ) [4]. An equivalent set of transversity angles is sometimes used in polarization analyses [5]. The differential decay width depends on complex amplitudes  $A_{\lambda_1\lambda_2}$ , corresponding to the  $X$ -daughter helicity states  $\lambda_i$ .



**Figure 1:** Definition of the production and helicity angles in the sequential process  $ab \rightarrow X \rightarrow P_1P_2 \rightarrow (p_{11}p_{12})(p_{21}p_{22})$ . The three helicity angles include  $\theta_1$  and  $\theta_2$ , defined in the rest frame of the two daughters  $P_1$  and  $P_2$ , and  $\Phi$ , defined in the  $X$  frame as the angle between the two decay planes. The two production angles  $\theta^*$  and  $\Psi$  are defined in the  $X$  frame, where  $\Psi$  is the angle between the production plane and the average of the two decay planes.

In the case of a spin-zero  $B$ -meson decay, its daughter helicities are constrained to  $\lambda_1 = \lambda_2 = \lambda$ . Therefore we simplify amplitude notation as  $A_\lambda$ . Moreover, most  $B$ -decay polarization analyses are limited to the case when the spin of one of the  $B$ -meson daughters is 1. In that case, there are only three independent amplitudes corresponding to  $\lambda = 0$  or  $\pm 1$  [6], where the last two can be expressed in terms of parity-even and parity-odd amplitudes  $A_{\parallel,\perp} = (A_{+1} \pm A_{-1})/\sqrt{2}$ . The overall decay amplitude involves three complex terms proportional to the above amplitudes and the Wigner  $d$  functions of helicity angles. The exact angular dependence would depend on the quantum numbers of the  $B$ -meson daughters and of their decay products, and can be found in the literature [6,7]. The differential decay rate would involve six real quantities  $\alpha_i$ , including interference terms,

$$\frac{d\Gamma}{\Gamma d\cos\theta_1 d\cos\theta_2 d\Phi} = \sum_i \alpha_i f_i(\cos\theta_1, \cos\theta_2, \Phi), \quad (1)$$

where each  $f_i(\cos\theta_1, \cos\theta_2, \Phi)$  has unique angular dependence specific to particle quantum numbers, and the  $\alpha_i$  parameters are defined as:

$$\alpha_1 = \frac{|A_0|^2}{\sum |A_\lambda|^2} = f_L, \quad (2)$$

$$\alpha_2 = \frac{|A_\parallel|^2 + |A_\perp|^2}{\sum |A_\lambda|^2} = (1 - f_L), \quad (3)$$

$$\alpha_3 = \frac{|A_\parallel|^2 - |A_\perp|^2}{\sum |A_\lambda|^2} = (1 - f_L - 2f_\perp), \quad (4)$$

$$\alpha_4 = \frac{\Im m(A_{\perp} A_{\parallel}^*)}{\Sigma |A_{\lambda}|^2} = \sqrt{f_{\perp}(1-f_L-f_{\perp})} \sin(\phi_{\perp}-\phi_{\parallel}), \quad (5)$$

$$\alpha_5 = \frac{\Re e(A_{\parallel} A_0^*)}{\Sigma |A_{\lambda}|^2} = \sqrt{f_L(1-f_L-f_{\perp})} \cos(\phi_{\parallel}), \quad (6)$$

$$\alpha_6 = \frac{\Im m(A_{\perp} A_0^*)}{\Sigma |A_{\lambda}|^2} = \sqrt{f_{\perp} f_L} \sin(\phi_{\perp}), \quad (7)$$

where the amplitudes have been expressed with the help of polarization parameters  $f_L$ ,  $f_{\perp}$ ,  $\phi_{\parallel}$ , and  $\phi_{\perp}$  defined in Table 1. Note that the terms proportional to  $\Re e(A_{\perp} A_{\parallel}^*)$ ,  $\Im m(A_{\parallel} A_0^*)$ , and  $\Re e(A_{\perp} A_0^*)$  are absent in Eqs. (2-7). However, these terms may appear for some three-body decays of a  $B$ -meson daughter, see Ref. 7.

**Table 1:** Rate, polarization, and  $CP$ -asymmetry parameters defined for the  $B$ -meson decays to mesons with non-zero spin. Numerical examples are shown for the average of the  $B^0 \rightarrow \varphi K^*(892)^0$  decay measurements obtained from BABAR [8], Belle [9], and LHCb [10]. The first six parameters are defined under the assumption of no  $CP$  violation in decay, while they are averaged between the  $\bar{B}$  and  $B$  parameters in general. The last six parameters involve differences between the  $\bar{B}$  and  $B$  meson decay parameters. The phase convention  $\delta_0$  is chosen with respect to a single  $A_{00}$  amplitude from a reference  $B$  decay mode, which is  $B^0 \rightarrow \varphi K_0^*(1430)^0$  for numerical results.

parameter	definition	average
$\mathcal{B}$	$\Gamma/\Gamma_{\text{total}}$	$(10.1^{+0.6}_{-0.5}) \times 10^{-6}$
$f_L$	$ A_0 ^2/\Sigma  A_{\lambda} ^2$	$0.497 \pm 0.017$
$f_{\perp}$	$ A_{\perp} ^2/\Sigma  A_{\lambda} ^2$	$0.225 \pm 0.015$
$\phi_{\parallel} - \pi$	$\arg(A_{\parallel}/A_0) - \pi$	$-0.712 \pm 0.058$
$\phi_{\perp} - \pi$	$\arg(A_{\perp}/A_0) - \pi$	$-0.615 \pm 0.056$
$\delta_0 - \pi$	$\arg(A_{00}/A_0) - \pi$	$-0.26 \pm 0.10$
$A_{CP}$	$(\bar{\Gamma} - \Gamma)/(\bar{\Gamma} + \Gamma)$	$-0.003 \pm 0.038$
$A_{CP}^0$	$(\bar{f}_L - f_L)/(\bar{f}_L + f_L)$	$-0.007 \pm 0.030$
$A_{CP}^{\perp}$	$(\bar{f}_{\perp} - f_{\perp})/(\bar{f}_{\perp} + f_{\perp})$	$-0.014 \pm 0.057$
$\Delta\phi_{\parallel}$	$(\bar{\phi}_{\parallel} - \phi_{\parallel})/2$	$+0.051 \pm 0.053$
$\Delta\phi_{\perp}$	$(\bar{\phi}_{\perp} - \phi_{\perp} - \pi)/2$	$+0.075 \pm 0.050$
$\Delta\delta_0$	$(\bar{\delta}_0 - \delta_0)/2$	$+0.13 \pm 0.08$

Overall, six real parameters describe three complex amplitudes  $A_0$ ,  $A_{\parallel}$ , and  $A_{\perp}$ . These could be chosen to be the four polarization parameters  $f_L$ ,  $f_{\perp}$ ,  $\phi_{\parallel}$ , and  $\phi_{\perp}$ , one overall size normalization, such as decay rate  $\Gamma$ , or branching fraction  $\mathcal{B}$ , and one overall phase  $\delta_0$ . The phase convention is arbitrary for an isolated  $B$  decay mode. However, for several  $B$  decays, the relative phase could produce meaningful and observable effects through interference with other  $B$  decays with the same final states, such as for  $B \rightarrow VK_J^*$  with  $J = 0, 1, 2, 3, 4, \dots$ . The phase could be referenced to the single  $B \rightarrow VK_0^*$  amplitude  $A_{00}$  in such a case, as shown in Table 1. Here  $V$  stands for any spin-one vector meson.

Moreover,  $CP$  violation can be tested in the angular distribution of the decay as the difference between the  $B$  and  $\bar{B}$ . Each of the six real parameters describing the three complex amplitudes would have a counterpart  $CP$ -asymmetry term, corresponding to three direct- $CP$  asymmetries in three amplitudes, and three  $CP$ -violating phase differences, equivalent to the phase measurements from the mixing-induced  $CP$  asymmetries in the time evolution of  $B$ -decays [1]. In Table 1 and Ref. 11, these are chosen to be the direct- $CP$  asymmetries in the overall decay rate  $\mathcal{A}_{CP}$ , in the  $f_L$  fraction  $\mathcal{A}_{CP}^0$ , and in the  $f_{\perp}$  fraction  $\mathcal{A}_{CP}^{\perp}$ , and three weak phase differences:

$$\Delta\phi_{\parallel} = \frac{1}{2} \arg(\bar{A}_{\parallel} A_0 / A_{\parallel} \bar{A}_0), \quad (8)$$

$$\Delta\phi_{\perp} = \frac{1}{2} \arg(\bar{A}_{\perp} A_0 / A_{\perp} \bar{A}_0) - \frac{\pi}{2}, \quad (9)$$

$$\Delta\delta_0 = \frac{1}{2} \arg(\bar{A}_{00} A_0 / A_{00} \bar{A}_0). \quad (10)$$

The  $\frac{\pi}{2}$  term in Eq. (9) reflects the fact that  $A_{\perp}$  and  $\bar{A}_{\perp}$  differ in phase by  $\pi$  if  $CP$  is conserved. The two parameters  $\Delta\phi_{\parallel}$  and  $\Delta\phi_{\perp}$  are equivalent to triple-product asymmetries constructed from the vectors describing the decay angular distribution [12]. The  $CP$ -violating phase difference in the reference decay mode [11] is, in the Wolfenstein CKM quark-mixing phase convention,

$$\Delta\phi_{00} = \frac{1}{2} \arg(A_{00}/\bar{A}_{00}). \quad (11)$$

This can be measured only together with the mixing-induced phase difference for some of the neutral  $B$ -meson decays similar to other mixing-induced  $CP$  asymmetry measurements [1].

It may not always be possible to have a phase-reference decay mode which would define  $\delta_0$  and  $\Delta\delta_0$  parameters. In that case, it may be possible to define the phase difference directly similarly to Eq. (11):

$$\Delta\phi_0 = \frac{1}{2} \arg(A_0/\bar{A}_0). \quad (12)$$

One can measure the angles of the CKM unitarity triangle, assuming Standard Model contributions to the  $\Delta\phi_0$  and  $B$ -mixing phases. Examples include measurements of  $\beta = \phi_1$  with  $B \rightarrow J/\psi K^*$  and  $\alpha = \phi_2$  with  $B \rightarrow \rho\rho$ .

Most of the  $B$  decays that arise from tree-level  $b \rightarrow c$  transitions have the amplitude hierarchy  $|A_0| > |A_{+}| > |A_{-}|$  which is expected from analyses based on quark-helicity conservation [13]. The larger the mass of the vector-meson daughters, the weaker the inequality. The  $B$  meson decays to heavy vector particles with charm, such as  $B \rightarrow J/\psi K^*$ ,  $\psi(2S)K^*$ ,  $\chi_{c1}K^*$ ,  $D^*\rho$ ,  $D^*K^*$ ,  $D^*D^*$ , and  $D^*D_s^*$ , show a substantial fraction of the amplitudes corresponding to transverse polarization of the vector mesons ( $A_{\pm 1}$ ), in agreement with the factorization prediction. The detailed amplitude analysis of the  $B \rightarrow J/\psi K^*$  decays has been performed by the BABAR [14], Belle [15], CDF [16], CLEO [17], D0 [18], and LHCb [19] collaborations. Most analyses are performed under the assumption of the absence of direct  $CP$  violation. The parameter values are given in

# Meson Particle Listings

## $B^0$

the particle listing of this *Review*. The difference between the strong phases  $\phi_{\parallel}$  and  $\phi_{\perp}$  deviates significantly from zero. The recent measurements [14,15] of  $CP$ -violating terms similar to those in  $B \rightarrow \varphi K^*$  [11] shown in Table 1 are consistent with zero.

In addition, the mixing-induced  $CP$ -violating asymmetry is measured in the  $B^0 \rightarrow J/\psi K^{*0}$  decay [1,14,15] where angular analysis allows one to separate  $CP$ -eigenstate amplitudes. This allows one to resolve the sign ambiguity of the  $\cos 2\beta$  ( $\cos 2\phi_1$ ) term that appears in the time-dependent angular distribution due to interference of parity-even and parity-odd terms. This analysis relies on the knowledge of discrete ambiguities in the strong phases  $\phi_{\parallel}$  and  $\phi_{\perp}$ , as discussed below. The BABAR experiment used a method based on the dependence on the  $K\pi$  invariant mass of the interference between the  $S$ - and  $P$ -waves to resolve the discrete ambiguity in the determination of the strong phases ( $\phi_{\parallel}, \phi_{\perp}$ ) in  $B \rightarrow J/\psi K^*$  decays [14]. The result is in agreement with the amplitude hierarchy expectation [13]. The CDF [20], D0 [21], and LHCb [22,23] experiments have studied the  $B_s^0 \rightarrow J/\psi \varphi$ ,  $J/\psi(K^+K^-)$ ,  $J/\psi(\pi^+\pi^-)$  decays and provided the lifetime, polarization, and phase measurements.

The amplitude hierarchy  $|A_0| \gg |A_+| \gg |A_-|$  was expected in  $B$  decays to light vector particles in both penguin transitions [24,25] and tree-level transitions [13]. There is confirmation by the BABAR and Belle experiments of predominantly longitudinal polarization in the tree-level  $b \rightarrow u$  transition, such as  $B^0 \rightarrow \rho^+\rho^-$  [26],  $B^+ \rightarrow \rho^0\rho^+$  [27], and  $B^+ \rightarrow \omega\rho^+$  [28]; this is consistent with the analysis of the quark helicity conservation [13]. Because the longitudinal amplitude dominates the decay, a detailed amplitude analysis is not possible with current  $B$  samples, and limits on the transverse amplitude fraction are obtained. The small branching fractions of  $B^0 \rightarrow \rho^0\rho^0, \omega\rho^0, \omega\omega$  [30,31,32,28] indicate that  $b \rightarrow d$  penguin pollution is small in the charmless, strangeless vector-vector  $B$  decays. There is a measurement of large longitudinal polarization in  $B^0 \rightarrow \rho^0\rho^0$  [30,31,32] decays. The fraction of transverse polarization is large in decays to heavier mesons such as  $B^0 \rightarrow a_1(1260)^+a_1(1260)^-$  [29].

The interest in the polarization and  $CP$ -asymmetry measurements in penguin transition, such as  $b \rightarrow s$  decays  $B \rightarrow \varphi K^*$ ,  $\rho K^*$ ,  $\omega K^*$ , or  $B_s^0 \rightarrow \varphi\varphi$ ,  $K^*K^*$ , and  $b \rightarrow d$  decay  $B \rightarrow K^*\bar{K}^*$ , is motivated by their potential sensitivity to physics beyond the Standard Model. The decay amplitudes for  $B \rightarrow \varphi K^*$  have been measured by the BABAR, Belle, and LHCb experiments [11,9,33,34,10]. The fractions of longitudinal polarization are  $f_L = 0.50 \pm 0.05$  for the  $B^+ \rightarrow \varphi K^{*+}$  decay and  $f_L = 0.497 \pm 0.017$  for the  $B^0 \rightarrow \varphi K^{*0}$  decay. These indicate significant departure from the naive expectation of predominant longitudinal polarization, suggesting other contributions to the decay amplitude, previously neglected, either within the Standard Model, such as penguin annihilation [35] or QCD rescattering [36], or from physics beyond the Standard Model [37]. The complete set of twelve amplitude parameters measured in the  $B^0 \rightarrow \varphi K^{*0}$  decay is given in Table 1. Several

other parameters could be constructed from the above twelve parameters, as suggested in Ref. 38.

The discrete ambiguity in the phase ( $\phi_{\parallel}, \phi_{\perp}, \Delta\phi_{\parallel}, \Delta\phi_{\perp}$ ) measurements has been resolved by BABAR in favor of  $|A_+| \gg |A_-|$  through interference between the  $S$ - and  $P$ -waves of  $K\pi$ . The search for vector-tensor and vector-axialvector  $B \rightarrow \varphi K_J^{(*)}$  decays with  $J = 1, 2, 3, 4$  revealed a large fraction of longitudinal polarization in the decay  $B \rightarrow \varphi K_2^*(1430)$  with  $f_L = 0.90_{-0.07}^{+0.06}$  [11,39], but large contribution of transverse amplitude in  $B \rightarrow \varphi K_1(1270)$  with  $f_L = 0.46_{-0.15}^{+0.13}$  [40].

Like  $B \rightarrow \varphi K^*$ , the decays  $B \rightarrow \rho K^*$  and  $B \rightarrow \omega K^*$  may be sensitive to New Physics. Measurements of the longitudinal polarization fraction in  $B^+ \rightarrow \rho^0 K^{*0}$ ,  $B^+ \rightarrow \rho^+ K^{*0}$  [41] and in both vector-vector and vector-tensor final states of  $B \rightarrow \omega K_J^*$  [28] reveal a large fraction of transverse polarization, indicating an anomaly similar to  $B \rightarrow \varphi K^*$  except for a different pattern in vector-tensor final states. A large transverse polarization is also observed in the  $B_s^0 \rightarrow \varphi\varphi$  decay by CDF [42] and LHCb [43],  $B_s^0 \rightarrow K^{*0}\bar{K}^{*0}$  decays by LHCb [44], and  $B_s^0 \rightarrow \varphi K^{*0}$  decays by LHCb [45]. At the same time, measurement of the polarization in the  $b \rightarrow d$  penguin decays  $B \rightarrow K^*\bar{K}^*$  indicates a large fraction of longitudinal polarization [46]. The polarization pattern in penguin-dominated  $B$ -meson decays is not fully understood [35,36,37].

The three-body semileptonic  $B$ -meson decays, such as  $B \rightarrow V\ell_1\ell_2$ , share many features with the two-body  $B \rightarrow VV$  decays. Their differential decay width can be parameterized with the two helicity angles defined in the  $V$  and  $(\ell_1\ell_2)$  frames and with the azimuthal angle, as defined in Fig. 1. However, since the  $(\ell_1\ell_2)$  pair does not come from an on-shell particle, the angular distribution is unique to each point in the dilepton mass  $m_{\ell\ell}$  spectrum. The polarization measurements as a function of  $m_{\ell\ell}$  provide complementary information on physics beyond the Standard Model, as discussed for  $B \rightarrow K^*\ell^+\ell^-$  and  $B_s \rightarrow \phi\ell^+\ell^-$  decays in Ref. 47. The current data in these modes have been analyzed by the BABAR, Belle, CDF, LHCb, and CMS experiments [48,49].

The examples of the angular distributions and observables in  $B \rightarrow K^*\ell^+\ell^-$  are discussed in Ref. 47. Typically two angular observables have been measured in this decay in certain ranges of the dilepton mass  $m_{\ell\ell}$  [48]. One parameter is the fraction of longitudinal polarization  $F_L$ , which is determined by the  $K^*$  angular distribution and is similar to  $f_L$  defined for exclusive two-body decays. The other parameter is the forward-backward asymmetry of the lepton pair  $A_{FB}$ , which is the asymmetry of the decay rate with positive and negative values of  $\cos\theta_1$ .

In summary, there has been considerable recent interest in the polarization measurements of  $B$ -meson decays because they reveal both weak- and strong-interaction dynamics [35–37,50]. New measurements will further elucidate the pattern of spin alignment measurements in rare  $B$  decays, and further test the Standard Model and strong interaction dynamics, including the non-factorizable contributions to the  $B$ -decay amplitudes.

See key on page 601

## References

1. M. Kreps and Y. Kwon, "Production and Decay of  $b$ -Flavored Hadrons," mini-review in this *Review*.
2. For a recent example and further references see Y.Y. Gao *et al.*, Phys. Rev. **D81**, 075022 (2010).
3. M. Jacob and G. C. Wick, Ann. Phys. **7**, 404 (1959).
4. S. M. Berman and M. Jacob, Phys. Rev. **139**, 1023 (1965).
5. I. Dunietz *et al.*, Phys. Rev. **D43**, 2193 (1991).
6. G. Kramer and W.F. Palmer, Phys. Rev. **D45**, 193 (1992).
7. A. Datta *et al.*, Phys. Rev. **D77**, 114025 (2008).
8. BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D78**, 092008 (2008).
9. Belle Collab., M. Prim *et al.*, Phys. Rev. **D88**, 072004 (2013).
10. LHCb Collab., R. Aaij *et al.*, JHEP **05**, 069 (2014).
11. BABAR Collab., B. Aubert *et al.*, Phys. Rev. Lett. **93**, 231804 (2004); Phys. Rev. Lett. **98**, 051801 (2007); Phys. Rev. **D78**, 092008 (2008).
12. G. Valencia, Phys. Rev. **D39**, 3339 (1998); A. Datta and D. London, Int. J. Mod. Phys. **A19**, 2505 (2004).
13. A. Ali *et al.*, Z. Phys. **C1**, 269 (1979); M. Suzuki, Phys. Rev. **D64**, 117503 (2001).
14. BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D71**, 032005 (2005); Phys. Rev. **D76**, 031102 (2007).
15. Belle Collab., R. Itoh *et al.*, Phys. Rev. Lett. **95**, 091601 (2005) Phys. Rev. **D88**, 072004 (2013).
16. CDF Collab., T. Affolder *et al.*, Phys. Rev. Lett. **85**, 4668 (2000); CDF Collab., D. Acosta *et al.*, Phys. Rev. Lett. **94**, 101803 (2005).
17. CLEO Collab., C. P. Jessop, Phys. Rev. Lett. **79**, 4533 (1997).
18. D0 Collab., V. M. Abazov *et al.*, Phys. Rev. Lett. **102**, 032001 (2009).
19. LHCb Collab., R. Aaij *et al.*, Phys. Rev. **D88**, 052002 (2013).
20. CDF Collab., T. Aaltonen *et al.*, Phys. Rev. Lett. **100**, 121803 (2008); Phys. Rev. **D85**, 072002 (2012).
21. D0 Collab., V. M. Abazov *et al.*, Phys. Rev. Lett. **98**, 121801 (2007); Phys. Rev. **D85**, 032006 (2012).
22. LHCb Collab., R. Aaij *et al.*, Phys. Rev. Lett. **108**, 101803 (2012).
23. LHCb Collab., R. Aaij *et al.*, Phys. Rev. Lett. **114**, 041801 (2015).
24. H.Y. Cheng and K.C. Yang, Phys. Lett. **B511**, 40 (2001); C.H. Chen, Y.Y. Keum, and H.n. Li, Phys. Rev. **D66**, 054013 (2002).
25. A.L.Kagan, Phys. Lett. **B601**, 151 (2004); Y. Grossman, Int. J. Mod. Phys. **A19**, 907 (2004).
26. Belle Collab., A. Somov *et al.*, Phys. Rev. Lett. **96**, 171801 (2006); BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D76**, 052007 (2007).
27. Belle Collab., J. Zhang *et al.*, Phys. Rev. Lett. **91**, 221801 (2003); BABAR Collab., B. Aubert *et al.*, Phys. Rev. Lett. **102**, 141802 (2009).
28. BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D74**, 051102 (2006); Phys. Rev. **D79**, 052005 (2009).
29. BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D80**, 092007 (2009).
30. BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D78**, 071104 (2008).
31. Belle Collab., I. Adachi *et al.*, Phys. Rev. **D89**, 072008 (2014).
32. LHCb Collab., R. Aaij *et al.*, Phys. Lett. **B747**, 468 (2015).
33. Belle Collab., K.F. Chen *et al.*, Phys. Rev. Lett. **94**, 221804 (2005).
34. BABAR Collab., B. Aubert *et al.*, Phys. Rev. Lett. **99**, 201802 (2007).
35. A.L. Kagan, Phys. Lett. **B601**, 151 (2004); H.n. Li and S. Mishima, Phys. Rev. **D71**, 054025 (2005); C.-H. Chen *et al.*, Phys. Rev. **D72**, 054011 (2005); M. Beneke *et al.*, Phys. Rev. Lett. **96**, 141801 (2006); C.-H. Chen and C.-Q. Geng, Phys. Rev. **D75**, 054010 (2007); A. Datta *et al.*, Phys. Rev. **D76**, 034015 (2007); M. Beneke, J. Rohrer, and D. Yang, Nucl. Phys. **B774**, 64 (2007); H.-Y. Cheng and K.-C. Yang, Phys. Rev. **D78**, 094001 (2008).
36. C. W. Bauer *et al.*, Phys. Rev. **D70**, 054015 (2004); P. Colangelo *et al.*, Phys. Lett. **B597**, 291 (2004); M. Ladisa *et al.*, Phys. Rev. **D70**, 114025 (2004); H. Y. Cheng *et al.*, Phys. Rev. **D71**, 014030 (2005); H. Y. Cheng and K. C. Yang, Phys. Rev. **D83**, 034001 (2011).
37. Y. Grossman, Int. J. Mod. Phys. **A19**, 907 (2004); E. Alvarez *et al.*, Phys. Rev. **D70**, 115014 (2004); P.K. Das and K.C. Yang, Phys. Rev. **D71**, 094002 (2005); C.H. Chen and C.Q. Geng, Phys. Rev. **D71**, 115004 (2005); Y.D. Yang *et al.*, Phys. Rev. **D72**, 015009 (2005); K.C. Yang, Phys. Rev. **72**, 034009 (2005); S. Baek, Phys. Rev. **D72**, 094008 (2005); C.S. Huang *et al.*, Phys. Rev. **D73**, 034026 (2006); C.H. Chen and H. Hatanaka, Phys. Rev. **D73**, 075003 (2006); A. Faessler *et al.*, Phys. Rev. **D75**, 074029 (2007).
38. D. London, N. Sinha, and R. Sinha, Phys. Rev. **D69**, 11401 (2004).
39. BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D76**, 051103 (2007).
40. BABAR Collab., B. Aubert *et al.*, Phys. Rev. Lett. **101**, 161801 (2008).
41. Belle Collab., J. Zhang *et al.*, Phys. Rev. Lett. **95**, 141801 (2005); BABAR Collab., B. Aubert *et al.*, Phys. Rev. Lett. **97**, 201801 (2006); Phys. Rev. **D83**, 051101 (2011); Phys. Rev. **D85**, 072005 (2012).
42. CDF Collab., T. Aaltonen *et al.*, Phys. Rev. Lett. **107**, 261802 (2011).
43. LHCb Collab., R. Aaij *et al.*, Phys. Lett. **B713**, 369 (2012); Phys. Rev. **D90**, 052011 (2014).
44. LHCb Collab., R. Aaij *et al.*, Phys. Lett. **B709**, 50 (2012); JHEP **1507**, 166 (2015).
45. LHCb Collab., R. Aaij *et al.*, JHEP **1311**, 092 (2013).
46. BABAR Collab., B. Aubert *et al.*, Phys. Rev. Lett. **100**, 081801 (2008), Phys. Rev. **D79**, 051102 (2009).
47. G. Burdman, Phys. Rev. **D52**, 6400 (1995); F. Kruger and J. Matias, Phys. Rev. **D71**, 094009 (2005); E. Lunghi and J. Matias, JHEP **0704**, 058 (2007); J. Matias *et al.*, JHEP **1204**, 104 (2012).
48. Belle Collab., J.-T. Wei *et al.*, Phys. Rev. Lett. **103**, 171801 (2009); BABAR Collab., B. Aubert *et al.*, Phys. Rev. **D79**, 031102 (2009); CDF Collab., T. Aaltonen *et al.*, Phys. Rev. Lett. **108**, 081807 (2012); LHCb Collab., R.

# Meson Particle Listings

$B^0$

Aaij *et al.*, JHEP **1308**, 131 (2013); JHEP **04**, 064 (2015); CMS Collab., S. Chatrchyan *et al.*, Phys. Lett. **B727**, 77 (2013).

49. LHCb Collab., R. Aaij *et al.*, JHEP **1509**, 179 (2015).
50. C.H. Chen and H.n. Li, Phys. Rev. **D71**, 114008 (2005).

## POLARIZATION IN $B^0$ DECAY

In decays involving two vector mesons, one can distinguish among the states in which meson polarizations are both longitudinal ( $L$ ) or both are transverse and parallel ( $\parallel$ ) or perpendicular ( $\perp$ ) to each other with the parameters  $\Gamma_L/\Gamma$ ,  $\Gamma_\parallel/\Gamma$ , and the relative phases  $\phi_\parallel$  and  $\phi_\perp$ . See the definitions in the note on "Polarization in  $B$  Decays" review in the  $B^0$  Particle Listings.

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow J/\psi(1S)K^*(892)^0$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.571 \pm 0.007</math> OUR AVERAGE</b>				
$0.572 \pm 0.006 \pm 0.014$		<sup>1</sup> AAIJ	13AT LHCb	$pp$ at 7 TeV
$0.587 \pm 0.011 \pm 0.013$		<sup>2</sup> ABAZOV	09E D0	$p\bar{p}$ at 1.96 TeV
$0.556 \pm 0.009 \pm 0.010$		<sup>3</sup> AUBERT	07AD BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.562 \pm 0.026 \pm 0.018$		ACOSTA	05 CDF	$p\bar{p}$ at 1.96 TeV
$0.574 \pm 0.012 \pm 0.009$		ITOH	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.59 \pm 0.06 \pm 0.01$		<sup>4</sup> AFFOLDER	00N CDF	$p\bar{p}$ at 1.8 TeV
$0.52 \pm 0.07 \pm 0.04$		<sup>5</sup> JESSOP	97 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$0.65 \pm 0.10 \pm 0.04$	65	ABE	95Z CDF	$p\bar{p}$ at 1.8 TeV
$0.97 \pm 0.16 \pm 0.15$	13	<sup>6</sup> ALBRECHT	94G ARG	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.566 \pm 0.012 \pm 0.005$		<sup>3</sup> AUBERT	05P BABR	Repl. by AUBERT 07AD
$0.62 \pm 0.02 \pm 0.03$		<sup>7</sup> ABE	02N BELL	Repl. by ITOH 05
$0.597 \pm 0.028 \pm 0.024$		<sup>8</sup> AUBERT	01H BABR	Repl. by AUBERT 07AD
$0.80 \pm 0.08 \pm 0.05$	42	<sup>6</sup> ALAM	94 CLE2	Sup. by JESSOP 97

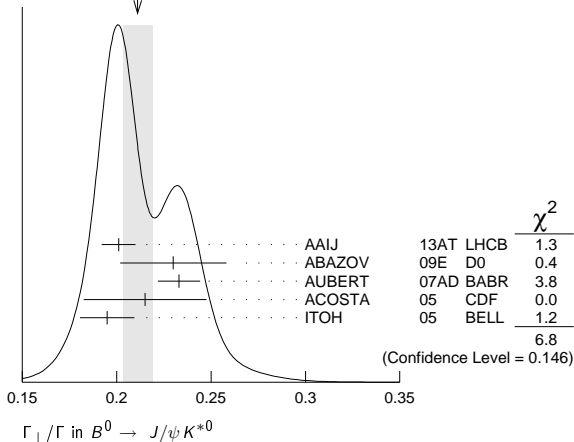
- <sup>1</sup> AAIJ 13AT obtains  $\Gamma_\parallel/\Gamma = 0.227 \pm 0.004 \pm 0.011$ . The relation  $1 = (\Gamma_L + \Gamma_\perp + \Gamma_\parallel)/\Gamma$  is used to obtain  $\Gamma_L/\Gamma$ .
- <sup>2</sup> Measured the angular and lifetime parameters for the time-dependent angular untagged decays  $B_d^0 \rightarrow J/\psi K^{*0}$  and  $B_s^0 \rightarrow J/\psi \phi$ .
- <sup>3</sup> Obtained by combining the  $B^0$  and  $B^+$  modes.
- <sup>4</sup> AFFOLDER 00N measurements are based on 190  $B^0$  candidates obtained from a data sample of  $89 \text{ pb}^{-1}$ . The  $P$ -wave fraction is found to be  $0.13^{+0.12}_{-0.09} \pm 0.06$ .
- <sup>5</sup> JESSOP 97 is the average over a mixture of  $B^0$  and  $B^+$  decays. The  $P$ -wave fraction is found to be  $0.16 \pm 0.08 \pm 0.04$ .
- <sup>6</sup> Averaged over an admixture of  $B^0$  and  $B^+$  decays.
- <sup>7</sup> Averaged over an admixture of  $B^0$  and  $B^+$  decays and the  $P$  wave fraction is  $(19 \pm 2 \pm 3)\%$ .
- <sup>8</sup> Averaged over an admixture of  $B^0$  and  $B^-$  decays and the  $P$  wave fraction is  $(16.0 \pm 3.2 \pm 1.4) \times 10^{-2}$ .

### $\Gamma_\perp/\Gamma$ in $B^0 \rightarrow J/\psi K^{*0}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.211 \pm 0.008</math> OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.		
$0.201 \pm 0.004 \pm 0.008$	AAIJ	13AT LHCb	$pp$ at 7 TeV
$0.230 \pm 0.013 \pm 0.025$	<sup>1</sup> ABAZOV	09E D0	$p\bar{p}$ at 1.96 TeV
$0.233 \pm 0.010 \pm 0.005$	<sup>2</sup> AUBERT	07AD BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.215 \pm 0.032 \pm 0.006$	ACOSTA	05 CDF	$p\bar{p}$ at 1.96 TeV
$0.195 \pm 0.012 \pm 0.008$	ITOH	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Measured the angular and lifetime parameters for the time-dependent angular untagged decays  $B_d^0 \rightarrow J/\psi K^{*0}$  and  $B_s^0 \rightarrow J/\psi \phi$ .
- <sup>2</sup> Obtained by combining the  $B^0$  and  $B^+$  modes.

WEIGHTED AVERAGE  
 $0.211 \pm 0.008$  (Error scaled by 1.3)

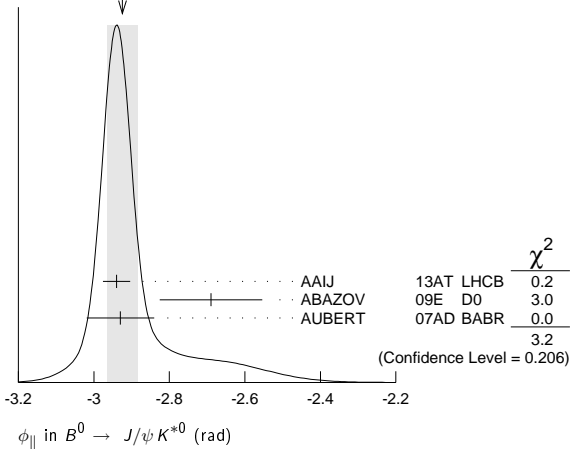


### $\phi_\parallel$ in $B^0 \rightarrow J/\psi K^{*0}$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>-2.92 \pm 0.04</math> OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.		
$-2.94 \pm 0.02 \pm 0.03$	AAIJ	13AT LHCb	$pp$ at 7 TeV
$-2.69 \pm 0.08 \pm 0.11$	<sup>1</sup> ABAZOV	09E D0	$p\bar{p}$ at 1.96 TeV
$-2.93 \pm 0.08 \pm 0.04$	<sup>2</sup> AUBERT	07AD BABR	$e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Obtained  $\phi_\parallel$  as  $\delta_2 - \delta_1$ , assuming they are uncorrelated.
- <sup>2</sup> Obtained by combining the  $B^0$  and  $B^+$  modes.

WEIGHTED AVERAGE  
 $-2.92 \pm 0.04$  (Error scaled by 1.3)

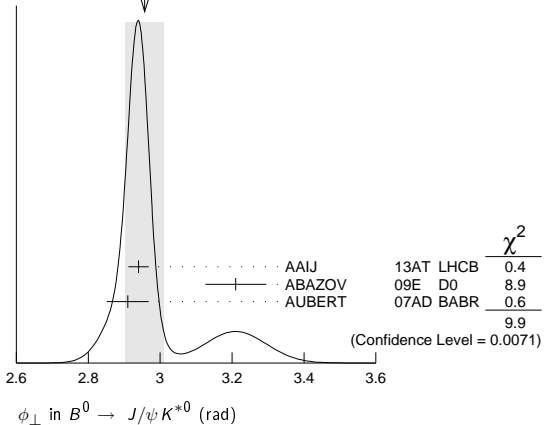


### $\phi_\perp$ in $B^0 \rightarrow J/\psi K^{*0}$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>2.96 \pm 0.05</math> OUR AVERAGE</b>	Error includes scale factor of 2.2. See the ideogram below.		
$2.94 \pm 0.02 \pm 0.02$	AAIJ	13AT LHCb	$pp$ at 7 TeV
$3.21 \pm 0.06 \pm 0.06$	ABAZOV	09E D0	$p\bar{p}$ at 1.96 TeV
$2.91 \pm 0.05 \pm 0.03$	<sup>1</sup> AUBERT	07AD BABR	$e^+e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Obtained by combining the  $B^0$  and  $B^+$  modes.

WEIGHTED AVERAGE  
 $2.96 \pm 0.05$  (Error scaled by 2.2)



### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \psi(2S)K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.463 \pm 0.028</math> OUR AVERAGE</b>			
$0.455 \pm 0.031 \pm 0.014$	CHILIKIN	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.48 \pm 0.05 \pm 0.02$	<sup>1</sup> AUBERT	07AD BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.45 \pm 0.11 \pm 0.04$	<sup>2</sup> RICHICHI	01 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.448 \pm 0.040 \pm 0.040$	MIZUK	09 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.027 \pm 0.053$			

- <sup>1</sup> Obtained by combining the  $B^0$  and  $B^+$  modes.
- <sup>2</sup> Averages between charged and neutral  $B$  mesons.



See key on page 601

# Meson Particle Listings

## $B^0$

### $\Gamma_{\perp}/\Gamma$ in $B^0 \rightarrow \psi(2S) K^{*0}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.30 \pm 0.06 \pm 0.02</math></b>	<sup>1</sup> AUBERT	07AD BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by combining the $B^0$ and $B^+$ modes.			

### $\phi_{\parallel}$ in $B^0 \rightarrow \psi(2S) K^{*0}$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>-2.8 \pm 0.4 \pm 0.1</math></b>	<sup>1</sup> AUBERT	07AD BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by combining the $B^0$ and $B^+$ modes.			

### $\phi_{\perp}$ in $B^0 \rightarrow \psi(2S) K^{*0}$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>2.8 \pm 0.3 \pm 0.1</math></b>	<sup>1</sup> AUBERT	07AD BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by combining the $B^0$ and $B^+$ modes.			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \chi_{c1} K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.83^{+0.06}_{-0.08}</math> OUR AVERAGE</b>	Error includes scale factor of 1.3.		
$0.947^{+0.038+0.046}_{-0.048-0.099}$	MIZUK	08 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.77 \pm 0.07 \pm 0.04$	<sup>1</sup> AUBERT	07AD BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by combining the $B^0$ and $B^+$ modes.			

### $\Gamma_{\perp}/\Gamma$ in $B^0 \rightarrow \chi_{c1} K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.03 \pm 0.04 \pm 0.02</math></b>	<sup>1</sup> AUBERT	07AD BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by combining the $B^0$ and $B^+$ modes.			

### $\phi_{\parallel}$ in $B^0 \rightarrow \chi_{c1} K^*(892)^0$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>0.0 \pm 0.3 \pm 0.1</math></b>	<sup>1</sup> AUBERT	07AD BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by combining the $B^0$ and $B^+$ modes.			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow D_s^{*+} D^{*-}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.52 \pm 0.05</math> OUR AVERAGE</b>			
$0.519 \pm 0.050 \pm 0.028$	<sup>1</sup> AUBERT	03i BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.506 \pm 0.139 \pm 0.036$	AHMED	00B CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Measurement performed using partial reconstruction of $D^{*-}$ decay.			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow D^{*-} \rho^+$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.885 \pm 0.016 \pm 0.012</math></b>		CSORNA	03 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.93 \pm 0.05 \pm 0.05$	76	ALAM	94 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow D_s^{*+} \rho^-$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.84^{+0.26}_{-0.28} \pm 0.13</math></b>	<sup>1</sup> AUBERT	08AJ BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow D_s^{*+} K^{*-}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.92^{+0.37}_{-0.31} \pm 0.07</math></b>	<sup>1</sup> AUBERT	08AJ BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow D^{*+} D^{*-}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.624 \pm 0.029 \pm 0.011</math></b>	KRONENBIT...12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.57 \pm 0.08 \pm 0.02$	MIYAKE	05 BELL	Repl. by KRONENBITTER 12

### $\Gamma_{\perp}/\Gamma$ in $B^0 \rightarrow D^{*+} D^{*-}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.147 \pm 0.019</math> OUR AVERAGE</b>			
$0.138 \pm 0.024 \pm 0.006$	KRONENBIT...12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.158 \pm 0.028 \pm 0.006$	AUBERT	09c BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.125 \pm 0.043 \pm 0.023$	VERVINK	09 BELL	Repl. by KRONENBITTER 12
$0.143 \pm 0.034 \pm 0.008$	AUBERT	07Bo BABR	Repl. by AUBERT 09c
$0.125 \pm 0.044 \pm 0.007$	AUBERT,BE	05A BABR	Repl. by AUBERT 07Bo
$0.19 \pm 0.08 \pm 0.01$	MIYAKE	05 BELL	Repl. by VERVINK 09
$0.063 \pm 0.055 \pm 0.009$	AUBERT	03Q BABR	Repl. by AUBERT,BE 05A

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \overline{D}^{*0} \omega$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.665 \pm 0.047 \pm 0.015</math></b>	LEES	11m BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \overline{D}_1(2430)^0 \omega$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>63.0 \pm 9.1^{+6.5}_{-6.0}</math></b>	<sup>1,2</sup> MATVIENKO	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by amplitude analysis of $\overline{B}^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.			
<sup>2</sup> Assumes equal production of $B^0$ and $B^+$ at $\Upsilon(4S)$ .			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \overline{D}_1(2420)^0 \omega$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>67.1 \pm 11.7^{+2.3}_{-5.0}</math></b>	<sup>1,2</sup> MATVIENKO	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by amplitude analysis of $\overline{B}^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.			
<sup>2</sup> Assumes equal production of $B^0$ and $B^+$ at $\Upsilon(4S)$ .			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \overline{D}_2^*(2460)^0 \omega$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>76.0^{+18.3+3.5}_{-8.5-2.8}</math></b>	<sup>1,2</sup> MATVIENKO	15 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Obtained by amplitude analysis of $\overline{B}^0 \rightarrow D^{*-} \omega \pi^+$ . The second uncertainty combines in quadrature experimental systematic and model uncertainties.			
<sup>2</sup> Assumes equal production of $B^0$ and $B^+$ at $\Upsilon(4S)$ .			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow D^{*-} \omega \pi^+$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.654 \pm 0.042 \pm 0.016</math></b>	<sup>1</sup> AUBERT	06L BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Invariant mass of the $[\omega \pi]$ system is restricted in the region 1.1 and 1.9 GeV.			

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \omega K^{*0}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.69 \pm 0.13</math> OUR AVERAGE</b>			
$0.72 \pm 0.14 \pm 0.02$	AUBERT	09H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.56 \pm 0.29^{+0.18}_{-0.08}$	GOLDENZWE..08	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \omega K_2^*(1430)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.45 \pm 0.12 \pm 0.02</math></b>	AUBERT	09H BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow K^{*0} \overline{K}^{*0}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.80^{+0.10}_{-0.12} \pm 0.06</math></b>	AUBERT	08i BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.497 \pm 0.017</math> OUR AVERAGE</b>			
$0.497 \pm 0.019 \pm 0.015$	AAIJ	14AMLHCB	$pp$ at 7 TeV
$0.499 \pm 0.030 \pm 0.018$	PRIM	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.494 \pm 0.034 \pm 0.013$	AUBERT	08Bg BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.506 \pm 0.040 \pm 0.015$	AUBERT	07D BABR	Repl. by AUBERT 08Bg
$0.45 \pm 0.05 \pm 0.02$	CHEN	05A BELL	Repl. by PRIM 13
$0.52 \pm 0.05 \pm 0.02$	<sup>1</sup> AUBERT,B	04W BABR	Repl. by AUBERT 07D
$0.65 \pm 0.07 \pm 0.02$	AUBERT	03V BABR	Repl. by AUBERT,B 04W
$0.41 \pm 0.10 \pm 0.04$	CHEN	03B BELL	Repl. by CHEN 05A

<sup>1</sup> AUBERT,B 04W also measures the fraction of parity-odd transverse contribution  $f_{\perp} = 0.22 \pm 0.05 \pm 0.02$  and the phases of the parity-even and parity-odd transverse amplitudes relative to the longitudinal amplitude.

### $\Gamma_{\perp}/\Gamma$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.224 \pm 0.015</math> OUR AVERAGE</b>			
$0.221 \pm 0.016 \pm 0.013$	AAIJ	14AMLHCB	$pp$ at 7 TeV
$0.238 \pm 0.026 \pm 0.008$	PRIM	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.212 \pm 0.032 \pm 0.013$	AUBERT	08Bg BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.227 \pm 0.038 \pm 0.013$	AUBERT	07D BABR	Repl. by AUBERT 08Bg
$0.31^{+0.06}_{-0.05} \pm 0.02$	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
$0.22 \pm 0.05 \pm 0.02$	AUBERT,B	04W BABR	Repl. by AUBERT 07D
<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.			

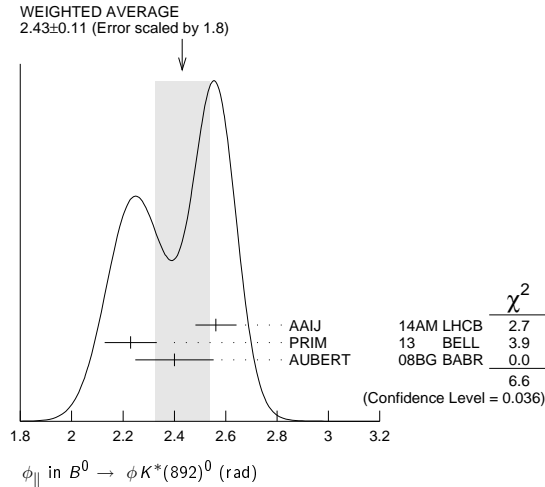
### $\phi_{\parallel}$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>2.43 \pm 0.11</math> OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.		
$2.562 \pm 0.069 \pm 0.040$	AAIJ	14AMLHCB	$pp$ at 7 TeV
$2.23 \pm 0.10 \pm 0.02$	PRIM	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$2.40 \pm 0.13 \pm 0.08$	AUBERT	08Bg BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.31 \pm 0.14 \pm 0.08$	AUBERT	07D BABR	Repl. by AUBERT 08Bg
$2.40^{+0.28}_{-0.24} \pm 0.07$	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
$2.34^{+0.23}_{-0.20} \pm 0.05$	AUBERT,B	04W BABR	Repl. by AUBERT 07D

# Meson Particle Listings

## $B^0$

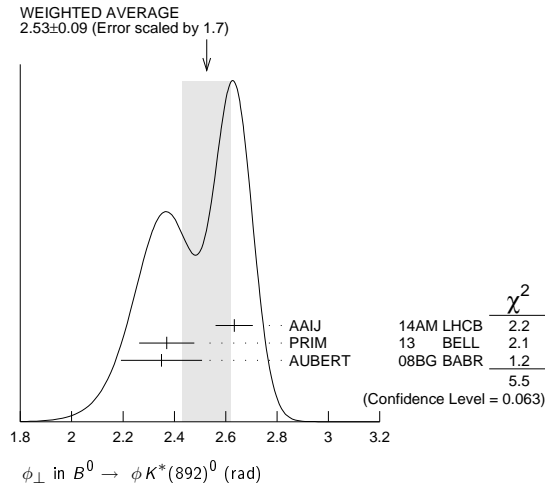
<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.



### $\phi_{\perp}$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>2.53 ± 0.09 OUR AVERAGE</b>	Error includes scale factor of 1.7. See the ideogram below.		
2.633 ± 0.062 ± 0.037	AAIJ	14AM LHCb	$pp$ at 7 TeV
2.37 ± 0.10 ± 0.04	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
2.35 ± 0.13 ± 0.09	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.24 ± 0.15 ± 0.09	AUBERT	07D BABR	Repl. by AUBERT 08Bg
2.51 ± 0.25 ± 0.06	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
2.47 ± 0.25 ± 0.05	AUBERT,B	04W BABR	Repl. by AUBERT 07D

<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.



### $\delta_0(B^0 \rightarrow \phi K^*(892)^0)$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>2.88 ± 0.10 OUR AVERAGE</b>			
2.91 ± 0.10 ± 0.08	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
2.82 ± 0.15 ± 0.09	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.78 ± 0.17 ± 0.09	AUBERT	07D BABR	Repl. by AUBERT 08Bg

### $A_{CP}^0$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.007 ± 0.030 OUR AVERAGE</b>			
-0.003 ± 0.038 ± 0.005	AAIJ	14AM LHCb	$pp$ at 7 TeV
-0.030 ± 0.061 ± 0.007	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.01 ± 0.07 ± 0.02	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.03 ± 0.08 ± 0.02	AUBERT	07D BABR	Repl. by AUBERT 08Bg
0.13 ± 0.12 ± 0.04	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
-0.06 ± 0.10 ± 0.01	AUBERT,B	04W BABR	Repl. by AUBERT 07D

<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.

### $A_{CP}^{\perp}$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.02 ± 0.06 OUR AVERAGE</b>			
0.047 ± 0.074 ± 0.009	AAIJ	14AM LHCb	$pp$ at 7 TeV
-0.14 ± 0.11 ± 0.01	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
-0.04 ± 0.15 ± 0.06	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.03 ± 0.16 ± 0.05	AUBERT	07D BABR	Repl. by AUBERT 08Bg
-0.20 ± 0.18 ± 0.04	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
-0.10 ± 0.24 ± 0.05	AUBERT,B	04W BABR	Repl. by AUBERT 07D

<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.

### $\Delta\phi_{\parallel}$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>0.05 ± 0.05 OUR AVERAGE</b>			
0.045 ± 0.062 ± 0.015	AAIJ	14AM LHCb	$pp$ at 7 TeV
-0.02 ± 0.10 ± 0.01	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.22 ± 0.12 ± 0.08	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.24 ± 0.14 ± 0.08	AUBERT	07D BABR	Repl. by AUBERT 08Bg
-0.32 ± 0.27 ± 0.07	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
0.27 ± 0.20 ± 0.05	AUBERT,B	04W BABR	Repl. by AUBERT 07D

<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.

### $\Delta\phi_{\perp}$ in $B^0 \rightarrow \phi K^*(892)^0$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>0.08 ± 0.05 OUR AVERAGE</b>			
0.062 ± 0.062 ± 0.005	AAIJ	14AM LHCb	$pp$ at 7 TeV
0.05 ± 0.10 ± 0.02	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.21 ± 0.13 ± 0.08	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.19 ± 0.15 ± 0.08	AUBERT	07D BABR	Repl. by AUBERT 08Bg
-0.30 ± 0.25 ± 0.06	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
0.36 ± 0.25 ± 0.05	AUBERT,B	04W BABR	Repl. by AUBERT 07D

<sup>1</sup> This quantity was recalculated by the BELLE authors from numbers in the original paper.

### $\Delta\delta_0(B^0 \rightarrow \phi K^*(892)^0)$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>0.13 ± 0.09 OUR AVERAGE</b>			
0.08 ± 0.10 ± 0.01	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.27 ± 0.14 ± 0.08	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.21 ± 0.17 ± 0.08	AUBERT	07D BABR	Repl. by AUBERT 08Bg

### $\Delta\phi_{00}(B^0 \rightarrow \phi K_0^*(1430)^0)$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>0.28 ± 0.42 ± 0.04</b>	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$

### $\Gamma_L/\Gamma$ in $B^0 \rightarrow \phi K_2^*(1430)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.913 ± 0.028 OUR AVERAGE</b>			
0.918 ± 0.029 ± 0.012	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.901 ± 0.046 ± 0.037	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.853 ± 0.061 ± 0.036	AUBERT	07D BABR	Repl. by AUBERT 08Bg

### $\Gamma_{\perp}/\Gamma$ in $B^0 \rightarrow \phi K_2^*(1430)^0$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.027 ± 0.031 OUR AVERAGE</b>	Error includes scale factor of 1.1.		
0.056 ± 0.050 ± 0.035	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.002 ± 0.018 ± 0.002	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.045 ± 0.049 ± 0.013	AUBERT	07D BABR	Repl. by AUBERT 08Bg

### $\phi_{\parallel}$ in $B^0 \rightarrow \phi K_2^*(1430)^0$

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b>4.0 ± 0.4 OUR AVERAGE</b>			
3.76 ± 2.88 ± 1.32	PRIM	13 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
3.96 ± 0.38 ± 0.06	AUBERT	08Bg BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.90 ± 0.39 ± 0.06	AUBERT	07D BABR	Repl. by AUBERT 08Bg

$\phi_L$  in  $B^0 \rightarrow \phi K_2^*(1430)^0$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
$4.45^{+0.43}_{-0.38} \pm 0.13$	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$5.72^{+0.55}_{-0.87} \pm 0.11$	AUBERT 07D	BABR	Repl. by AUBERT 08BG

 $\delta_0(B^0 \rightarrow \phi K_2^*(1430)^0)$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>3.46 \pm 0.14</math> OUR AVERAGE</b>			
$3.53 \pm 0.11 \pm 0.19$	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$3.41 \pm 0.13 \pm 0.13$	AUBERT 08BG	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.54^{+0.12}_{-0.14} \pm 0.06$	AUBERT 07D	BABR	Repl. by AUBERT 08BG

 $A_{CP}^0$  in  $B^0 \rightarrow \phi K_2^*(1430)^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.03 \pm 0.04</math> OUR AVERAGE</b>			
$-0.016^{+0.066}_{-0.051} \pm 0.008$	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.05 \pm 0.06 \pm 0.01$	AUBERT 08BG	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $A_{CP}^\perp$  in  $B^0 \rightarrow \phi K_2^*(1430)^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.01^{+0.85}_{-0.67} \pm 0.09$	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Delta\phi_{||}(B^0 \rightarrow \phi K_2^*(1430)^0)$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>-0.9 \pm 0.4</math> OUR AVERAGE</b>			
$-0.02 \pm 1.08 \pm 1.01$	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-1.00 \pm 0.38 \pm 0.09$	AUBERT 08BG	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Delta\phi_\perp(B^0 \rightarrow \phi K_2^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.19 \pm 0.42 \pm 0.11</math></b>	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Delta\delta_0$  in  $B^0 \rightarrow \phi K_2^*(1430)^0$ 

VALUE (rad)	DOCUMENT ID	TECN	COMMENT
<b><math>0.08 \pm 0.09</math> OUR AVERAGE</b>			
$0.06 \pm 0.11 \pm 0.02$	PRIM 13	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.11 \pm 0.13 \pm 0.06$	AUBERT 08BG	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow K^*(892)^0 \rho^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.40 \pm 0.08 \pm 0.11</math></b>	LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.57 \pm 0.09 \pm 0.08$	AUBERT,B 06G	BABR	Repl. by LEES 12K

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow K^{*+} \rho^-$ 

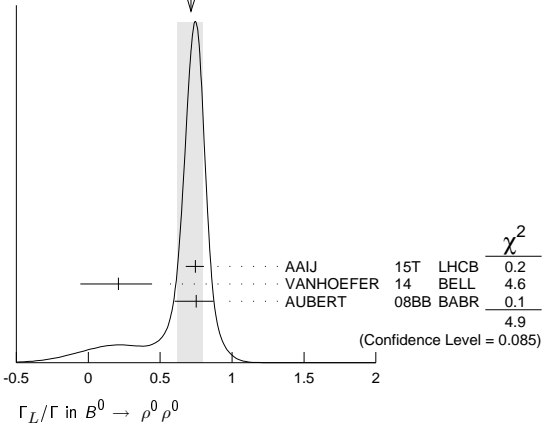
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.38 \pm 0.13 \pm 0.03</math></b>	LEES 12K	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow \rho^+ \rho^-$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.990^{+0.021}_{-0.019}</math> OUR AVERAGE</b>			
$0.988 \pm 0.012 \pm 0.023$	VANHOEFER 16	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.992 \pm 0.024^{+0.026}_{-0.013}$	AUBERT 07BF	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.941^{+0.034}_{-0.040} \pm 0.030$	SOMOV 06	BELL	Repl. by VANHOEFER 16
$0.978 \pm 0.014^{+0.021}_{-0.029}$	AUBERT,B 05C	BABR	Repl. by AUBERT 07BF
$0.98^{+0.02}_{-0.08} \pm 0.03$	AUBERT 04G	BABR	Repl. by AUBERT,B 04R
$0.99 \pm 0.03^{+0.04}_{-0.03}$	AUBERT,B 04R	BABR	Repl. by AUBERT,B 05C

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow \rho^0 \rho^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.71^{+0.08}_{-0.09}</math> OUR AVERAGE</b>			Error includes scale factor of 1.6. See the ideogram below.
$0.745^{+0.048}_{-0.058} \pm 0.034$	AAIJ 15T	LHCB	$pp$ at 7, 8 TeV
$0.21^{+0.18}_{-0.22} \pm 0.15$	VANHOEFER 14	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.75^{+0.11}_{-0.14} \pm 0.05$	AUBERT 08BB	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.87 \pm 0.13 \pm 0.04$	AUBERT 07G	BABR	Repl. by AUBERT 08BB

WEIGHTED AVERAGE  
0.71±0.08-0.09 (Error scaled by 1.6) $\Gamma_L/\Gamma$  in  $B^0 \rightarrow a_1(1260)^+ a_1(1260)^-$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.31 \pm 0.22 \pm 0.10</math></b>	AUBERT 09AL	BABR	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow \rho \bar{\rho} K^*(892)^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.01 \pm 0.13 \pm 0.03</math></b>	CHEN 08C	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow \Lambda \bar{\Lambda} K^*(892)^0$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.60 \pm 0.22 \pm 0.08</math></b>	CHANG 09	BELL	$e^+e^- \rightarrow \Upsilon(4S)$

 $\Gamma_L/\Gamma$  in  $B^0 \rightarrow K^*(892)^0 e^+ e^-$  ( $0.002 < q^2 < 1.120 \text{ GeV}^2/c^4$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.16 \pm 0.06 \pm 0.03</math></b>	AAIJ 15Z	LHCB	$pp$ at 7, 8 TeV

 $A_{\Gamma}^{(2)}$  in  $B^0 \rightarrow K^*(892)^0 e^+ e^-$  ( $0.002 < q^2 < 1.120 \text{ GeV}^2/c^4$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.23 \pm 0.23 \pm 0.05</math></b>	AAIJ 15Z	LHCB	$pp$ at 7, 8 TeV

 $A_{\Gamma}^{Im}$  in  $B^0 \rightarrow K^*(892)^0 e^+ e^-$  ( $0.002 < q^2 < 1.120 \text{ GeV}^2/c^4$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.14 \pm 0.22 \pm 0.05</math></b>	AAIJ 15Z	LHCB	$pp$ at 7, 8 TeV

 $A_{\Gamma}^{Re}$  in  $B^0 \rightarrow K^*(892)^0 e^+ e^-$  ( $0.002 < q^2 < 1.120 \text{ GeV}^2/c^4$ )

VALUE	DOCUMENT ID	TECN	COMMENT
Related to $A_{FB}$ , $F_L$ by $A_{\Gamma}^{Re} = (4/3) A_{FB} / (1 - F_L)$ .			
<b><math>0.10 \pm 0.18 \pm 0.05</math></b>	AAIJ 15Z	LHCB	$pp$ at 7, 8 TeV

 $B^0$ - $\bar{B}^0$  MIXING

Updated April 2016 by O. Schneider (Ecole Polytechnique Fédérale de Lausanne).

There are two neutral  $B^0$ - $\bar{B}^0$  meson systems,  $B_d^0$ - $\bar{B}_d^0$  and  $B_s^0$ - $\bar{B}_s^0$  (generically denoted  $B_q^0$ - $\bar{B}_q^0$ ,  $q = s, d$ ), which exhibit particle-antiparticle mixing [1]. This mixing phenomenon is described in Ref. 2. In the following, we adopt the notation introduced in Ref. 2, and assume  $CPT$  conservation throughout. In each system, the light (L) and heavy (H) mass eigenstates,

$$|B_{L,H}\rangle = p|B_q^0\rangle \pm q|\bar{B}_q^0\rangle, \quad (1)$$

have a mass difference  $\Delta m_q = m_H - m_L > 0$ , and a total decay width difference  $\Delta\Gamma_q = \Gamma_L - \Gamma_H$ . In the absence of  $CP$  violation in the mixing,  $|q/p| = 1$ , these differences are given by  $\Delta m_q = 2|M_{12}|$  and  $|\Delta\Gamma_q| = 2|\Gamma_{12}|$ , where  $M_{12}$  and  $\Gamma_{12}$  are the off-diagonal elements of the mass and decay matrices [2]. The evolution of a pure  $|B_q^0\rangle$  or  $|\bar{B}_q^0\rangle$  state at  $t = 0$  is given by

$$|B_q^0(t)\rangle = g_+(t)|B_q^0\rangle + \frac{q}{p}g_-(t)|\bar{B}_q^0\rangle, \quad (2)$$

$$|\bar{B}_q^0(t)\rangle = g_+(t)|\bar{B}_q^0\rangle + \frac{p}{q}g_-(t)|B_q^0\rangle, \quad (3)$$

# Meson Particle Listings

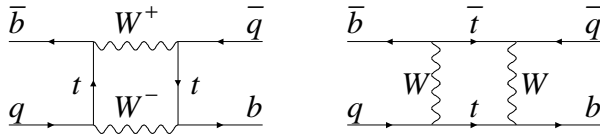
## $B^0$

which means that the flavor states remain unchanged (+) or oscillate into each other (−) with time-dependent probabilities proportional to

$$|g_{\pm}(t)|^2 = \frac{e^{-\Gamma_q t}}{2} \left[ \cosh\left(\frac{\Delta\Gamma_q}{2} t\right) \pm \cos(\Delta m_q t) \right], \quad (4)$$

where  $\Gamma_q = (\Gamma_H + \Gamma_L)/2$ . In the absence of  $CP$  violation, the time-integrated mixing probability  $\int |g_{-}(t)|^2 dt / (\int |g_{-}(t)|^2 dt + \int |g_{+}(t)|^2 dt)$  is given by

$$\chi_q = \frac{x_q^2 + y_q^2}{2(x_q^2 + 1)}, \quad \text{where} \quad x_q = \frac{\Delta m_q}{\Gamma_q}, \quad y_q = \frac{\Delta\Gamma_q}{2\Gamma_q}. \quad (5)$$



**Figure 1:** Dominant box diagrams for the  $B_q^0 \rightarrow \bar{B}_q^0$  transitions ( $q = d$  or  $s$ ). Similar diagrams exist where one or both  $t$  quarks are replaced with  $c$  or  $u$  quarks.

### Standard Model predictions and phenomenology

In the Standard Model, the transitions  $B_q^0 \rightarrow \bar{B}_q^0$  and  $\bar{B}_q^0 \rightarrow B_q^0$  are due to the weak interaction. They are described, at the lowest order, by box diagrams involving two  $W$  bosons and two up-type quarks (see Fig. 1), as is the case for  $K^0 - \bar{K}^0$  mixing. However, the long range interactions arising from intermediate virtual states are negligible for the neutral  $B$  meson systems, because the large  $B$  mass is off the region of hadronic resonances. The calculation of the dispersive and absorptive parts of the box diagrams yields the following predictions for the off-diagonal element of the mass and decay matrices [3],

$$M_{12} = -\frac{G_F^2 m_W^2 \eta_B m_{B_q} B_{B_q} f_{B_q}^2}{12\pi^2} S_0(m_t^2/m_W^2) (V_{tq}^* V_{tb})^2, \quad (6)$$

$$\Gamma_{12} = \frac{G_F^2 m_b^2 \eta_B' m_{B_q} B_{B_q} f_{B_q}^2}{8\pi} \times \left[ (V_{tq}^* V_{tb})^2 + V_{tq}^* V_{tb} V_{cq}^* V_{cb} \mathcal{O}\left(\frac{m_c^2}{m_b^2}\right) + (V_{cq}^* V_{cb})^2 \mathcal{O}\left(\frac{m_c^4}{m_b^4}\right) \right], \quad (7)$$

where  $G_F$  is the Fermi constant,  $m_W$  the  $W$  boson mass, and  $m_i$  the mass of quark  $i$ ;  $m_{B_q}$ ,  $f_{B_q}$  and  $B_{B_q}$  are the  $B_q^0$  mass, weak decay constant and bag parameter, respectively. The known function  $S_0(x_t)$  can be approximated very well by  $0.784 x_t^{0.76}$  [4], and  $V_{ij}$  are the elements of the CKM matrix [5]. The QCD corrections  $\eta_B$  and  $\eta_B'$  are of order unity. The only non-negligible contributions to  $M_{12}$  are from box diagrams involving two top quarks. The phases of  $M_{12}$  and  $\Gamma_{12}$  satisfy

$$\phi_M - \phi_\Gamma = \pi + \mathcal{O}\left(\frac{m_c^2}{m_b^2}\right), \quad (8)$$

implying that the mass eigenstates have mass and width differences of opposite signs. This means that, like in the  $K^0 - \bar{K}^0$  system, the heavy state is expected to have a smaller decay width than that of the light state:  $\Gamma_H < \Gamma_L$ . Hence,  $\Delta\Gamma = \Gamma_L - \Gamma_H$  is expected to be positive in the Standard Model.

Furthermore, the quantity

$$\left| \frac{\Gamma_{12}}{M_{12}} \right| \simeq \frac{3\pi}{2} \frac{m_b^2}{m_W^2} \frac{1}{S_0(m_t^2/m_W^2)} \sim \mathcal{O}\left(\frac{m_b^2}{m_t^2}\right) \quad (9)$$

is small, and a power expansion of  $|q/p|^2$  yields

$$\left| \frac{q}{p} \right|^2 = 1 + \left| \frac{\Gamma_{12}}{M_{12}} \right| \sin(\phi_M - \phi_\Gamma) + \mathcal{O}\left(\left| \frac{\Gamma_{12}}{M_{12}} \right|^2\right). \quad (10)$$

Therefore, considering both Eqs. (8) and (9), the  $CP$ -violating parameter

$$1 - \left| \frac{q}{p} \right|^2 \simeq \text{Im}\left(\frac{\Gamma_{12}}{M_{12}}\right) \quad (11)$$

is expected to be very small:  $\sim \mathcal{O}(10^{-3})$  for the  $B_d^0 - \bar{B}_d^0$  system and  $\lesssim \mathcal{O}(10^{-4})$  for the  $B_s^0 - \bar{B}_s^0$  system [6].

In the approximation of negligible  $CP$  violation in mixing, the ratio  $\Delta\Gamma_q/\Delta m_q$  is equal to the small quantity  $|\Gamma_{12}/M_{12}|$  of Eq. (9); it is hence independent of CKM matrix elements, *i.e.*, the same for the  $B_d^0 - \bar{B}_d^0$  and  $B_s^0 - \bar{B}_s^0$  systems. Calculations [7] yield  $\sim 5 \times 10^{-3}$  with a  $\sim 20\%$  uncertainty. Given the published experimental knowledge [8] on the mixing parameter  $x_q$

$$\begin{cases} x_d = 0.775 \pm 0.006 & (B_d^0 - \bar{B}_d^0 \text{ system}) \\ x_s = 26.81 \pm 0.10 & (B_s^0 - \bar{B}_s^0 \text{ system}) \end{cases}, \quad (12)$$

the Standard Model thus predicts that  $\Delta\Gamma_d/\Gamma_d$  is very small (below 1%), but  $\Delta\Gamma_s/\Gamma_s$  considerably larger ( $\sim 10\%$ ). These width differences are caused by the existence of final states to which both the  $B_q^0$  and  $\bar{B}_q^0$  mesons can decay. Such decays involve  $b \rightarrow c\bar{c}q$  quark-level transitions, which are Cabibbo-suppressed if  $q = d$  and Cabibbo-allowed if  $q = s$ .

A complete set of Standard Model predictions for all mixing parameters in both the  $B_d^0 - \bar{B}_d^0$  and  $B_s^0 - \bar{B}_s^0$  systems can be found in Ref. 9.

### Experimental issues and methods for oscillation analyses

Time-integrated measurements of  $B^0 - \bar{B}^0$  mixing were published for the first time in 1987 by UA1 [10] and ARGUS [11], and since then by many other experiments. These measurements are typically based on counting same-sign and opposite-sign lepton pairs from the semileptonic decay of the produced  $b\bar{b}$  pairs. Such analyses cannot easily separate the contributions from the different  $b$ -hadron species, therefore, the clean environment of  $\Upsilon(4S)$  machines (where only  $B_d^0$  and charged  $B_u$  mesons are produced) is in principle best suited to measure  $\chi_d$ .

However, better sensitivity is obtained from time-dependent analyses aiming at the direct measurement of the oscillation frequencies  $\Delta m_d$  and  $\Delta m_s$ , from the proper time distributions of

$B_d^0$  or  $B_s^0$  candidates identified through their decay in (mostly) flavor-specific modes, and suitably tagged as mixed or unmixed. This is particularly true for the  $B_s^0\text{--}\bar{B}_s^0$  system, where the large value of  $x_s$  implies maximal mixing, *i.e.*,  $\chi_s \simeq 1/2$ . In such analyses, the  $B_d^0$  or  $B_s^0$  mesons are either fully reconstructed, partially reconstructed from a charm meson, selected from a lepton with the characteristics of a  $b \rightarrow \ell^-$  decay, or selected from a reconstructed displaced vertex. At high-energy colliders (LEP, SLC, Tevatron, LHC), the proper time  $t = \frac{m_B}{p}L$  is measured from the distance  $L$  between the production vertex and the  $B$  decay vertex, and from an estimate of the  $B$  momentum  $p$ . At asymmetric  $B$  factories (KEKB, PEP-II), producing  $e^+e^- \rightarrow \Upsilon(4S) \rightarrow B_d^0\bar{B}_d^0$  events with a boost  $\beta\gamma$  ( $= 0.425, 0.55$ ), the proper time difference between the two  $B$  candidates is estimated as  $\Delta t \simeq \frac{\Delta z}{\beta\gamma c}$ , where  $\Delta z$  is the spatial separation between the two  $B$  decay vertices along the boost direction. In all cases, the good resolution needed on the vertex positions is obtained with silicon detectors.

The average statistical significance  $\mathcal{S}$  of a  $B_d^0$  or  $B_s^0$  oscillation signal can be approximated as [12]

$$\mathcal{S} \approx \sqrt{N/2} f_{\text{sig}} (1 - 2\eta) e^{-(\Delta m \sigma_t)^2/2}, \quad (13)$$

where  $N$  is the number of selected and tagged candidates,  $f_{\text{sig}}$  is the fraction of signal in that sample,  $\eta$  is the total mistag probability, and  $\sigma_t$  is the resolution on proper time (or proper time difference). The quantity  $\mathcal{S}$  decreases very quickly as  $\Delta m$  increases; this dependence is controlled by  $\sigma_t$ , which is therefore a critical parameter for  $\Delta m_s$  analyses. At high-energy colliders, the proper time resolution  $\sigma_t \sim \frac{m_B}{\langle p \rangle} \sigma_L \oplus t \frac{\sigma_p}{p}$  includes a constant contribution due to the decay length resolution  $\sigma_L$  (typically 0.04–0.3 ps), and a term due to the relative momentum resolution  $\sigma_p/p$  (typically 10–20% for partially reconstructed decays), which increases with proper time. At  $B$  factories, the boost of the  $B$  mesons is estimated from the known beam energies, and the term due to the spatial resolution dominates (typically 1–1.5 ps because of the much smaller  $B$  boost).

In order to tag a  $B$  candidate as mixed or unmixed, it is necessary to determine its flavor both in the initial state and in the final state. The initial and final state mistag probabilities,  $\eta_i$  and  $\eta_f$ , degrade  $\mathcal{S}$  by a total factor  $(1 - 2\eta) = (1 - 2\eta_i)(1 - 2\eta_f)$ . In lepton-based analyses, the final state is tagged by the charge of the lepton from  $b \rightarrow \ell^-$  decays; the largest contribution to  $\eta_f$  is then due to  $\bar{b} \rightarrow \bar{c} \rightarrow \ell^-$  decays. Alternatively, the charge of a reconstructed charm meson ( $D^{*-}$  from  $B_d^0$  or  $D_s^-$  from  $B_s^0$ ), or that of a kaon hypothesized to come from a  $b \rightarrow c \rightarrow s$  decay [13], can be used. For fully-inclusive analyses based on topological vertexing, final-state tagging techniques include jet-charge [14] and charge-dipole [15,16] methods. At high-energy colliders, the methods to tag the initial state (*i.e.*, the state at production), can be divided into two groups: the ones that tag the initial charge of the  $\bar{b}$  quark contained in the  $B$  candidate itself (same-side tag), and the ones that tag the initial charge of the other  $b$  quark produced in the event

(opposite-side tag). On the same side, the sign of a charged pion, kaon or proton from the primary vertex is correlated with the production state of the  $B_d^0$  or  $B_s^0$  if that particle is a decay product of a  $B^{**}$  state or the first in the fragmentation chain [17,18]. Jet- and vertex-charge techniques work on both sides and on the opposite side, respectively. Finally, the charge of a lepton from  $b \rightarrow \ell^-$ , of a kaon from  $b \rightarrow c \rightarrow s$  or of a charm hadron from  $b \rightarrow c$  [19] can be used as opposite side tags, keeping in mind that their performance is degraded due to integrated mixing. At SLC, the beam polarization produced a sizeable forward-backward asymmetry in the  $Z \rightarrow b\bar{b}$  decays, and provided another very interesting and effective initial state tag based on the polar angle of the  $B$  candidate [15]. Initial state tags have also been combined to reach  $\eta_i \sim 26\%$  at LEP [18,20], or even 22% at SLD [15] with full efficiency. In the case  $\eta_f = 0$ , this corresponds to an effective tagging efficiency  $Q = \epsilon D^2 = \epsilon(1 - 2\eta)^2$ , where  $\epsilon$  is the tagging efficiency, in the range 23–31%. The equivalent figure achieved by CDF during Tevatron Run I was  $\sim 3.5\%$  [21], reflecting the fact that tagging is more difficult at hadron colliders. The CDF and DØ analyses of Tevatron Run II data reached  $\epsilon D^2 = (1.8 \pm 0.1)\%$  [22] and  $(2.5 \pm 0.2)\%$  [23] for opposite-side tagging, while same-side kaon tagging (for  $B_s^0$  analyses) contributed an additional 3.7–4.8% at CDF [22], and pushed the combined performance to  $(4.7 \pm 0.5)\%$  at DØ [24]. LHCb, operating in the forward region at the LHC where the environment is different in terms of track multiplicity and  $b$ -hadron production kinematics, has reported  $\epsilon D^2 = (2.10 \pm 0.25)\%$  [25] for opposite-side tagging and  $(1.80 \pm 0.26)\%$  [26] for same-side kaon tagging, with a combined figure ranging typically between  $(3.73 \pm 0.15)\%$  [27] and  $(5.33 \pm 0.25)\%$  [28] depending on the mode in which the tagged meson is reconstructed.

At  $B$  factories, the flavor of a  $B_d^0$  meson at production cannot be determined, since the two neutral  $B$  mesons produced in a  $\Upsilon(4S)$  decay evolve in a coherent  $P$ -wave state where they keep opposite flavors at any time. However, as soon as one of them decays, the other follows a time-evolution given by Eqs. (2) or (3), where  $t$  is replaced with  $\Delta t$  (which will take negative values half of the time). Hence, the “initial state” tag of a  $B$  can be taken as the final-state tag of the other  $B$ . Effective tagging efficiencies of 30% are achieved by BaBar and Belle [29], using different techniques including  $b \rightarrow \ell^-$  and  $b \rightarrow c \rightarrow s$  tags. It is worth noting that, in this case, mixing of the other  $B$  (*i.e.*, the coherent mixing occurring before the first  $B$  decay) does not contribute to the mistag probability.

Before the experimental observation of a decay-width difference, oscillation analyses typically neglected  $\Delta\Gamma$  in Eq. (4), and described the physics with the functions  $\Gamma e^{-\Gamma t}(1 \pm \cos(\Delta m t))/2$  (high-energy colliders) or  $\Gamma e^{-\Gamma|\Delta t|}(1 \pm \cos(\Delta m \Delta t))/4$  (asymmetric  $\Upsilon(4S)$  machines). As can be seen from Eq. (4), a non-zero value of  $\Delta\Gamma$  would effectively reduce the oscillation amplitude with a small time-dependent factor that would be very difficult to distinguish from time resolution effects. Measurements

# Meson Particle Listings

## $B^0$

of  $\Delta m$  are usually extracted from the data using a maximum likelihood fit.

### $\Delta m_d$ and $\Delta \Gamma_d$ measurements

Many  $B_d^0$ - $\bar{B}_d^0$  oscillations analyses have been published [30] by the ALEPH [31], DELPHI [16,32], L3 [33], OPAL [34,35] BaBar [36], Belle [37], CDF [17], DØ [23], and LHCb [38–40] collaborations. Although a variety of different techniques have been used, the individual  $\Delta m_d$  results obtained at LEP and Tevatron have remarkably similar precision. Their average is compatible with the recent and more precise measurements from the asymmetric  $B$  factories and the LHC. The systematic uncertainties are not negligible; they are often dominated by sample composition, mistag probability, or  $b$ -hadron lifetime contributions. Before being combined, the measurements are adjusted on the basis of a common set of input values, including the  $b$ -hadron lifetimes and fractions published in this *Review*. Some measurements are statistically correlated. Systematic correlations arise both from common physics sources (fragmentation fractions, lifetimes, branching ratios of  $b$  hadrons), and from purely experimental or algorithmic effects (efficiency, resolution, tagging, background description). Combining all published [16,17,23,31–40] or recently submitted [41] measurements and accounting for all identified correlations yields  $\Delta m_d = 0.5065 \pm 0.0016(\text{stat}) \pm 0.0011(\text{syst}) \text{ ps}^{-1}$  [8], a result dominated by the new LHCb measurement with  $B^0 \rightarrow D^{(*)-} \mu^+ \nu_\mu X$  decays [41].

On the other hand, ARGUS and CLEO have published time-integrated measurements [42–44], which average to  $\chi_d = 0.182 \pm 0.015$ . Following Ref. 44, the width difference  $\Delta \Gamma_d$  could in principle be extracted from the measured value of  $\Gamma_d$  and the above averages for  $\Delta m_d$  and  $\chi_d$  (see Eq. (5)), provided that  $\Delta \Gamma_d$  has a negligible impact on the  $\Delta m_d$  measurements. However, direct time-dependent studies published by DELPHI [16], BaBar [45], Belle [46] and LHCb [47] provide stronger constraints, which can be combined to yield [8]

$$\Delta \Gamma_d / \Gamma_d = -0.003 \pm 0.015. \quad (14)$$

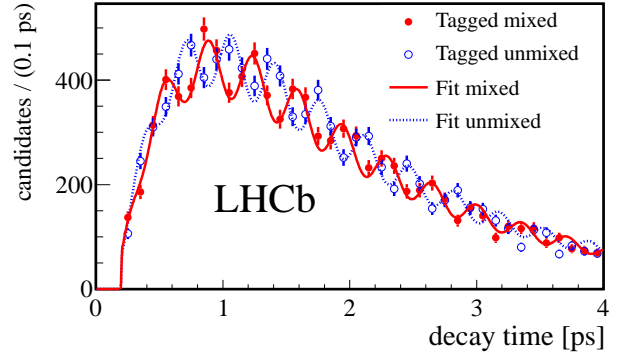
Assuming  $\Delta \Gamma_d = 0$  and no  $CP$  violation in mixing, and using the measured  $B_d^0$  lifetime of  $1.520 \pm 0.004 \text{ ps}$ , the  $\Delta m_d$  and  $\chi_d$  results are combined to yield the world average [48]

$$\Delta m_d = 0.5064 \pm 0.0019 \text{ ps}^{-1} \quad (15)$$

or, equivalently,

$$\chi_d = 0.1860 \pm 0.0011. \quad (16)$$

This  $\Delta m_d$  value provides an estimate of  $2|M_{12}|$ , and can be used with Eq. (6) to extract  $|V_{td}|$  within the Standard Model [49]. The main experimental uncertainties on the result come from  $m_t$  and  $\Delta m_d$ , but are completely negligible with respect to the uncertainty due to the hadronic matrix element  $f_{B_d} \sqrt{B_{B_d}} = 216 \pm 15 \text{ MeV}$  [50] obtained from unquenched lattice QCD calculations.



**Figure 2:** Proper time distribution of  $B_s^0 \rightarrow D_s^- \pi^+$  candidates tagged as mixed (red) or unmixed (blue) in the LHCb experiment, displaying  $B_s^0$ - $\bar{B}_s^0$  oscillations (from Ref. [51]).

### $\Delta m_s$ and $\Delta \Gamma_s$ measurements

After many years of intense search at LEP and SLC,  $B_s^0$ - $\bar{B}_s^0$  oscillations were first observed in 2006 by CDF using  $1 \text{ fb}^{-1}$  of Tevatron Run II data [22]. More recently LHCb observed  $B_s^0$ - $\bar{B}_s^0$  oscillations independently with  $B_s^0 \rightarrow D_s^- \pi^+$  [38,51],  $B_s^0 \rightarrow D_s^- \mu^+ \nu X$  [40] and even  $B_s^0 \rightarrow J/\psi K^+ K^-$  [27] decays, using between 1 and  $3 \text{ fb}^{-1}$  of data collected at the LHC until the end of 2012. Taking systematic correlations into account, the average of all published measurements of  $\Delta m_s$  [22,27,38,40,51] is

$$\Delta m_s = 17.757 \pm 0.020(\text{stat}) \pm 0.007(\text{syst}) \text{ ps}^{-1}, \quad (17)$$

dominated by LHCb (see Fig. 2) and still statistically limited.

The information on  $|V_{ts}|$  obtained in the framework of the Standard Model is hampered by the hadronic uncertainty, as in the  $B_d^0$  case. However, several uncertainties cancel in the frequency ratio

$$\frac{\Delta m_s}{\Delta m_d} = \frac{m_{B_s}}{m_{B_d}} \xi^2 \left| \frac{V_{ts}}{V_{td}} \right|^2, \quad (18)$$

where  $\xi = (f_{B_s} \sqrt{B_{B_s}}) / (f_{B_d} \sqrt{B_{B_d}}) = 1.268 \pm 0.063$  is an SU(3) flavor-symmetry breaking factor obtained from unquenched lattice QCD calculations [50]. Using the measurements of Eqs. (15) and (17), one can extract

$$\left| \frac{V_{td}}{V_{ts}} \right| = 0.2159 \pm 0.0004(\text{exp}) \pm 0.0107(\text{lattice}), \quad (19)$$

in good agreement with (but much more precise than) the value obtained from the ratio of the  $b \rightarrow d\gamma$  and  $b \rightarrow s\gamma$  transition rates observed at the  $B$  factories [49].

The CKM matrix can be constrained using experimental results on observables such as  $\Delta m_d$ ,  $\Delta m_s$ ,  $|V_{ub}/V_{cb}|$ ,  $\epsilon_K$ , and  $\sin(2\beta)$  together with theoretical inputs and unitarity conditions [49,52,53]. The constraint from our knowledge on the ratio  $\Delta m_s/\Delta m_d$  is more effective in limiting the position of the apex of the CKM unitarity triangle than the one obtained from the  $\Delta m_d$  measurements alone, due to the reduced hadronic uncertainty in Eq. (18). We also note that the measured value of  $\Delta m_s$  is consistent with the Standard Model prediction obtained

from CKM fits where no experimental information on  $\Delta m_s$  is used, *e.g.*,  $17.5 \pm 1.1 \text{ ps}^{-1}$  [52] or  $16.73^{+0.82}_{-0.57} \text{ ps}^{-1}$  [53].

Information on  $\Delta\Gamma_s$  can be obtained from the study of the proper time distribution of untagged  $B_s^0$  samples [54]. In the case of an inclusive  $B_s^0$  selection [55], or a flavor-specific (semileptonic or hadronic)  $B_s^0$  decay selection [20,56,57], both the short- and long-lived components are present, and the proper time distribution is a superposition of two exponentials with decay constants  $\Gamma_{L,H} = \Gamma_s \pm \Delta\Gamma_s/2$ . In principle, this provides sensitivity to both  $\Gamma_s$  and  $(\Delta\Gamma_s/\Gamma_s)^2$ . Ignoring  $\Delta\Gamma_s$  and fitting for a single exponential leads to an estimate of  $\Gamma_s$  with a relative bias proportional to  $(\Delta\Gamma_s/\Gamma_s)^2$ . An alternative approach, which is directly sensitive to first order in  $\Delta\Gamma_s/\Gamma_s$ , is to determine the effective lifetime of untagged  $B_s^0$  candidates decaying to pure  $CP$  eigenstates; measurements exist for  $B_s^0 \rightarrow K^+K^-$  [58],  $B_s^0 \rightarrow D_s^+D_s^-$  [57],  $B_s^0 \rightarrow J/\psi f_0(980)$  [59],  $B_s^0 \rightarrow J/\psi\pi^+\pi^-$  [60] and  $B_s^0 \rightarrow J/\psi K_S^0$  [61]. The extraction of  $1/\Gamma_s$  and  $\Delta\Gamma_s$  from such measurements, discussed in detail in Ref. [62], requires additional information in the form of theoretical assumptions or external inputs on weak phases and hadronic parameters. In what follows, the effective lifetimes from the above decays to pure  $CP$  eigenstates will be assumed to be dominated by a single weak phase.

The best sensitivity to  $1/\Gamma_s$  and  $\Delta\Gamma_s$  is achieved by the time-dependent measurements of the  $B_s^0 \rightarrow J/\psi\phi$  (or more generally  $B_s^0 \rightarrow J/\psi K^+K^-$ ) decay rates performed at CDF [63], DØ [64], ATLAS [65,66], CMS [67] and LHCb [27], where the  $CP$ -even and  $CP$ -odd amplitudes are separated statistically through a full angular analysis. The LHCb collaboration analyzes the  $B_s^0 \rightarrow J/\psi K^+K^-$  decay considering that the  $K^+K^-$  system can be in a P-wave or S-wave state, and measures the dependence of the strong phase difference between the P-wave and S-wave amplitudes as a function of the  $K^+K^-$  invariant mass [27,68]; this allows the unambiguous determination of the sign of  $\Delta\Gamma_s$ , which is found to be positive. All these studies use both untagged and tagged  $B_s^0$  candidates and are optimized for the measurement of the  $CP$ -violating phase  $\phi_s$ , defined as the weak phase difference between the  $B_s^0\text{--}\bar{B}_s^0$  mixing amplitude and the  $b \rightarrow c\bar{c}s$  decay amplitude. As reported below in Eq. (28), the current experimental average of  $\phi_s$  is consistent with zero. Assuming no  $CP$  violation (*i.e.*,  $\phi_s = 0$ ) a combination [8] of the published  $B_s^0 \rightarrow J/\psi\phi$ ,  $J/\psi K^+K^-$  analyses [27,63–65] and of the published effective lifetime measurements with flavor-specific [20,56,57] and pure  $CP$  [57–61] final states yields

$$\Delta\Gamma_s = +0.082 \pm 0.007 \text{ ps}^{-1} \quad \text{and} \quad 1/\Gamma_s = 1.510 \pm 0.005 \text{ ps}, \quad (20)$$

or, equivalently,

$$1/\Gamma_L = 1.422 \pm 0.008 \text{ ps} \quad \text{and} \quad 1/\Gamma_H = 1.610 \pm 0.012 \text{ ps}, \quad (21)$$

in good agreement with the Standard Model prediction  $\Delta\Gamma_s = 0.088 \pm 0.020 \text{ ps}^{-1}$  [9].

**Table 1:**  $\bar{\chi}$  and  $b$ -hadron fractions (see text).

	in $Z$ decays [8]	at Tevatron [8]	at LHC [80,82]
$\bar{\chi}$	$0.1259 \pm 0.0042$	$0.147 \pm 0.011$	
$f_u = f_d$	$0.407 \pm 0.007$	$0.344 \pm 0.021$	
$f_s$	$0.100 \pm 0.008$	$0.115 \pm 0.013$	
$f_{\text{baryon}}$	$0.085 \pm 0.011$	$0.197 \pm 0.046$	
$f_s/f_d$	$0.246 \pm 0.023$	$0.334 \pm 0.041$	$0.249 \pm 0.014$

Estimates of  $\Delta\Gamma_s/\Gamma_s$  obtained from measurements of the  $B_s^0 \rightarrow D_s^{(*)+}D_s^{(*)-}$  branching fractions are not included in the average, since they are based on the questionable [7] assumption that these decays account for all  $CP$ -even final states.

#### Average $b$ -hadron mixing probability and $b$ -hadron production fractions at high energy

Mixing measurements can significantly improve our knowledge on the fractions  $f_u$ ,  $f_d$ ,  $f_s$ , and  $f_{\text{baryon}}$ , defined as the fractions of  $B_u$ ,  $B_d$ ,  $B_s^0$ , and  $b$ -baryons in an unbiased sample of weakly decaying  $b$  hadrons produced in high-energy collisions. Indeed, time-integrated mixing analyses using lepton pairs from  $b\bar{b}$  events at high energy measure the quantity

$$\bar{\chi} = f'_d \chi_d + f'_s \chi_s, \quad (22)$$

where  $f'_d$  and  $f'_s$  are the fractions of  $B_d^0$  and  $B_s^0$  hadrons in a sample of semileptonic  $b$ -hadron decays. Assuming that all  $b$  hadrons have the same semileptonic decay width implies  $f'_q = f_q/(\Gamma_q\tau_b)$  ( $q = s, d$ ), where  $\tau_b$  is the average  $b$ -hadron lifetime. Hence  $\bar{\chi}$  measurements performed at LEP [69] and Tevatron [70,71], together with the  $\chi_d$  average of Eq. (16) and the very good approximation  $\chi_s = 1/2$  (in fact  $\chi_s = 0.499308 \pm 0.000005$  from Eqs. (5), (17) and (20)), provide constraints on the fractions  $f_d$  and  $f_s$ .

The LEP experiments have measured  $\mathcal{B}(\bar{b} \rightarrow B_s^0) \times \mathcal{B}(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell X)$  [72],  $\mathcal{B}(b \rightarrow \Lambda_b^0) \times \mathcal{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell X)$  [73], and  $\mathcal{B}(b \rightarrow \Xi_b^-) \times \mathcal{B}(\Xi_b^- \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X)$  [74] from partially reconstructed final states including a lepton,  $f_{\text{baryon}}$  from protons identified in  $b$  events [75], and the production rate of charged  $b$  hadrons [76]. The  $b$ -hadron fraction ratios measured at CDF are based on double semileptonic  $K^*\mu\mu$  and  $\phi\mu\mu$  final states [77] and lepton-charm final states [78]; in addition CDF and DØ have both measured strange  $b$ -baryon production [79]. On the other hand, fraction ratios have been studied by LHCb using fully reconstructed hadronic  $B_s^0$  and  $B_d^0$  decays [80], as well as semileptonic decays [81]. ATLAS has measured  $f_s/f_d$  using  $B_s^0 \rightarrow J/\psi\phi$  and  $B^0 \rightarrow J/\psi K^{*0}$  decays [82]. Both CDF and LHCb observe that the ratio  $f_{\Lambda_b^0}/(f_u + f_d)$  decreases with the transverse momentum of the lepton+charm system, indicating that the  $b$ -hadron fractions are not the same in different environments. We therefore provide sets of fractions separately for LEP and Tevatron (and no complete set for LHC, where strange  $b$ -baryon production has not been measured yet). A combination of all the available information under the constraints  $f_u = f_d$ ,  $f_u + f_d + f_s + f_{\text{baryon}} = 1$ , and Eq. (22), yields the averages shown in the first two columns of Table 1.

# Meson Particle Listings

## $B^0$

### $CP$ -violation studies

Evidence for  $CP$  violation in  $B_q^0\text{--}\bar{B}_q^0$  mixing has been searched for, both with flavor-specific and inclusive  $B_q^0$  decays, in samples where the initial flavor state is tagged, usually with a lepton from the other  $b$ -hadron in the event. In the case of semileptonic (or other flavor-specific) decays, where the final-state tag is also available, the following asymmetry [2]

$$\mathcal{A}_{\text{SL}}^q = \frac{N(\bar{B}_q^0(t) \rightarrow \ell^+ \nu_\ell X) - N(B_q^0(t) \rightarrow \ell^- \bar{\nu}_\ell X)}{N(\bar{B}_q^0(t) \rightarrow \ell^+ \nu_\ell X) + N(B_q^0(t) \rightarrow \ell^- \bar{\nu}_\ell X)} \simeq 1 - |q/p|_q^2 \quad (23)$$

has been measured either in time-integrated analyses at CLEO [44,83], BaBar [84], CDF [85], DØ [86–88] and LHCb [89], or in time-dependent analyses at LEP [35,90], BaBar [45,91] and Belle [92]. In the inclusive case, also investigated at LEP [90,93], no final-state tag is used, and the asymmetry [94]

$$\begin{aligned} & \frac{N(\bar{B}_q^0(t) \rightarrow \text{all}) - N(B_q^0(t) \rightarrow \text{all})}{N(\bar{B}_q^0(t) \rightarrow \text{all}) + N(B_q^0(t) \rightarrow \text{all})} \\ & \simeq \mathcal{A}_{\text{SL}}^q \left[ \sin^2 \left( \frac{\Delta m_q t}{2} \right) - \frac{x_q}{2} \sin(\Delta m_q t) \right] \end{aligned} \quad (24)$$

must be measured as a function of the proper time to extract information on  $CP$  violation. In addition LHCb has studied the time dependence of the charge asymmetry of  $B^0 \rightarrow D^{(*)-} \mu^+ \nu_\mu X$  decays without tagging the initial state [95], which would be equal to

$$\frac{N(D^{(*)-} \mu^+ \nu_\mu X) - N(D^{(*)+} \mu^- \bar{\nu}_\mu X)}{N(D^{(*)-} \mu^+ \nu_\mu X) + N(D^{(*)+} \mu^- \bar{\nu}_\mu X)} = \mathcal{A}_{\text{SL}}^d \frac{1 - \cos(\Delta m_d t)}{2} \quad (25)$$

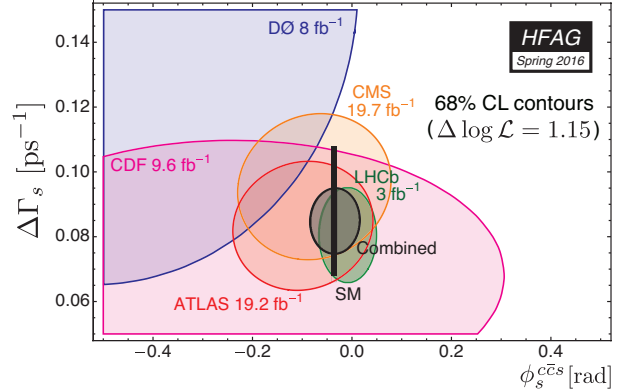
in absence of detection and production asymmetries.

The DØ collaboration measures a like-sign dimuon charge asymmetry in semileptonic  $b$  decays that deviates by  $2.8\sigma$  from the tiny Standard Model prediction and concludes, from a more refined analysis in bins of muon impact parameters, that the overall discrepancy is at the level of  $3.6\sigma$  [86]. In all other cases, asymmetries compatible with zero (and the Standard Model [9]) have been found, with a precision limited by the available statistics. Several of the analyses at high energy don't disentangle the  $B_d^0$  and  $B_s^0$  contributions, and either quote a mean asymmetry or a measurement of  $\mathcal{A}_{\text{SL}}^d$  assuming  $\mathcal{A}_{\text{SL}}^s = 0$ : we no longer include these in the average. An exception is the latest dimuon DØ analysis [86], which separates the two contributions by exploiting their dependence on the muon impact parameter cut. The resulting measurements of  $\mathcal{A}_{\text{SL}}^d$  and  $\mathcal{A}_{\text{SL}}^s$  are then both compatible with the Standard Model. They are also correlated. We therefore perform a two-dimensional average of the measurements of Refs. [44,45,83,84,86–89,91,92,95] and obtain [8]

$$\mathcal{A}_{\text{SL}}^d = -0.0015 \pm 0.0017, \text{ or } |q/p|_d = 1.0007 \pm 0.0009, \quad (26)$$

$$\mathcal{A}_{\text{SL}}^s = -0.0075 \pm 0.0041, \text{ or } |q/p|_s = 1.0038 \pm 0.0021, \quad (27)$$

with a correlation coefficient of  $-0.16$  between  $\mathcal{A}_{\text{SL}}^d$  and  $\mathcal{A}_{\text{SL}}^s$ . These results show no evidence of  $CP$  violation and don't constrain yet the Standard Model.



**Figure 3:** 68% CL contours in the  $(\phi_s, \Delta\Gamma_s)$  plane, showing the measurements from CDF [63], DØ [64], ATLAS [65,66], CMS [67] and LHCb [27,28,96], with their combination [8]. The thin rectangle represents the Standard Model predictions of  $\phi_s$  [53] and  $\Delta\Gamma_s$  [9].

$CP$  violation induced by  $B_s^0\text{--}\bar{B}_s^0$  mixing in  $b \rightarrow c\bar{c}s$  decays has been a field of very active study in the past few years. In addition to the previously mentioned  $B_s^0 \rightarrow J/\psi\phi$  and  $B_s^0 \rightarrow J/\psi K^+ K^-$  studies, the decay modes  $B_s^0 \rightarrow J/\psi\pi^+\pi^-$  (including  $B_s^0 \rightarrow J/\psi f_0(980)$ ) [96] and  $B_s^0 \rightarrow D_s^+ D_s^-$  [28] have also been analyzed by LHCb to measure  $\phi_s$ , without the need for an angular analysis. The  $J/\psi\pi^+\pi^-$  final state has been shown indeed to be (very close to) a pure  $CP$ -odd state [97]. A two-dimensional fit [8] of all these results [27,28,63–67,96] in the  $(\phi_s, \Delta\Gamma_s)$  plane, shown on Fig. 3, yields [48]

$$\phi_s = -0.033 \pm 0.033. \quad (28)$$

This is consistent with the Standard Model prediction for  $\phi_s$ , which is equal to  $-2\beta_s = -2\arg(-(V_{ts}V_{tb}^*)/(V_{cs}V_{cb}^*)) = -0.0376^{+0.0008}_{-0.0007}$  [53], assuming negligible Penguin pollution.

### Summary

$B^0\text{--}\bar{B}^0$  mixing has been and still is a field of intense study. The mass differences in the  $B_d^0\text{--}\bar{B}_d^0$  and  $B_s^0\text{--}\bar{B}_s^0$  systems are now known to relative precisions of 0.38% and 0.12%, respectively. The non-zero decay width difference in the  $B_s^0\text{--}\bar{B}_s^0$  system is well established, with a relative difference of  $\Delta\Gamma_s/\Gamma_s = (12.4 \pm 1.1)\%$ , meaning that the heavy state of the  $B_s^0\text{--}\bar{B}_s^0$  system lives  $\sim 13\%$  longer than the light state. In contrast, the relative decay width difference in the  $B_d^0\text{--}\bar{B}_d^0$  system,  $\Delta\Gamma_d/\Gamma_d = (-0.3 \pm 1.5)\%$ , is still consistent with zero.  $CP$  violation in mixing has not been observed yet, with precisions on the semileptonic asymmetries below 0.5%. An impressive progress has been achieved in the measurement of the mixing-induced phase  $\phi_s$  in  $B_s^0$  decays proceeding through the  $b \rightarrow c\bar{c}s$  transition, with an uncertainty of 33 mrad. Despite these significant improvements, all observations remain consistent with the Standard Model expectations.

However, the measurements where New Physics might show up are still statistically limited. More results are awaited from the LHC experiments and Belle II, with promising prospects for the investigation of the  $CP$ -violating phase  $\arg(-M_{12}/\Gamma_{12})$  and an improved determination of  $\phi_s$ .



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Mixing studies have clearly reached the stage of precision measurements, where much effort is needed, both on the experimental and theoretical sides, in particular to further reduce the hadronic uncertainties of lattice QCD calculations. In the long term, a stringent check of the consistency of the  $B_d^0$  and  $B_s^0$  mixing amplitudes (magnitudes and phases) with all other measured flavor-physics observables will be possible within the Standard Model, leading to very tight limits on (or otherwise a long-awaited surprise about) New Physics.

## References

1. T.D. Lee and C.S. Wu, *Ann. Rev. Nucl. Sci.* **16**, 511 (1966); I.I. Bigi and A.I. Sanda, “*CP* violation,” Cambridge Univ. Press, 2000; G.C. Branco, L. Lavoura, and J.P. Silva, “*CP* violation,” Clarendon Press Oxford, 1999.
2. See the review on *CP* violation in the quark sector by T. Gershon and Y. Nir in this publication.
3. A.J. Buras, W. Slominski, and H. Steger, *Nucl. Phys.* **B245**, 369 (1984).
4. T. Inami and C.S. Lim, *Prog. Theor. Phys.* **65**, 297 (1981); for the power-like approximation, see A.J. Buras and R. Fleischer, page 91 in “Heavy Flavours II,” eds. A.J. Buras and M. Lindner, Singapore World Scientific, 1998.
5. M. Kobayashi and K. Maskawa, *Prog. Theor. Phys.* **49**, 652 (1973).
6. I.I. Bigi *et al.*, in “*CP* violation,” ed. C. Jarlskog, Singapore World Scientific, 1989.
7. A. Lenz and U. Nierste, [arXiv:1102.4274 \[hep-ph\]](https://arxiv.org/abs/1102.4274); A. Lenz and U. Nierste, *JHEP* **06**, 072 (2007).
8. Y. Amhis *et al.* (HFAG), “Averages of *b*-hadron, *c*-hadron, and  $\tau$ -lepton properties as of summer 2014,” [arXiv:1412.7515 \[hep-ex\]](https://arxiv.org/abs/1412.7515); the combined results on *b*-hadron fractions, lifetimes and mixing parameters published in this *Review* have been obtained by the *B* oscillations working group of the Heavy Flavor Averaging Group (HFAG), using the methods and procedures described in Chapter 3 of the above paper, after updating the list of inputs; for more information, see <http://www.slac.stanford.edu/xorg/hfag/osc/>.
9. M. Artuso, G. Borissov, and A. Lenz, [arXiv:1511.09466 \[hep-ph\]](https://arxiv.org/abs/1511.09466).
10. C. Albajar *et al.* [UA1 Collab.], *Phys. Lett.* **B186**, 247 (1987).
11. H. Albrecht *et al.* [ARGUS Collab.], *Phys. Lett.* **B192**, 245 (1987).
12. H.-G. Moser and A. Roussarie, *Nucl. Instrum. Methods* **A384**, 491 (1997).
13. SLD Collab., SLAC-PUB-7228, SLAC-PUB-7229, and SLAC-PUB-7230, *28th Int. Conf. on High Energy Physics*, Warsaw, 1996; J. Wittlin, PhD thesis, SLAC-R-582, 2001.
14. ALEPH Collab., contrib. 596 to *Int. Europhysics Conf. on High Energy Physics*, Jerusalem, 1997.
15. K. Abe *et al.* [SLD Collab.], *Phys. Rev.* **D67**, 012006 (2003).
16. J. Abdallah *et al.* [DELPHI Collab.], *Eur. Phys. J.* **C28**, 155 (2003).
17. F. Abe *et al.* [CDF Collab.], *Phys. Rev. Lett.* **80**, 2057 (1998) and *Phys. Rev.* **D59**, 032001 (1999); *Phys. Rev.* **D60**, 051101 (1999); *Phys. Rev.* **D60**, 072003 (1999); T. Affolder *et al.* [CDF Collab.], *Phys. Rev.* **D60**, 112004 (1999).
18. R. Barate *et al.* [ALEPH Collab.], *Eur. Phys. J.* **C4**, 367 (1998); *Eur. Phys. J.* **C7**, 553 (1999).
19. R. Aaij *et al.* [LHCb Collab.], *JINST* **10**, P10005 (2015).
20. P. Abreu *et al.* [DELPHI Collab.], *Eur. Phys. J.* **C16**, 555 (2000).
21. See tagging summary on page 160 of K. Anikeev *et al.*, “*B* physics at the Tevatron: Run II and beyond,” FERMILAB-PUB-01/97, [hep-ph/0201071](https://arxiv.org/abs/hep-ph/0201071), and references therein.
22. A. Abulencia *et al.* [CDF Collab.], *Phys. Rev. Lett.* **97**, 242003 (2006).
23. V.M. Abazov *et al.* [DØ Collab.], *Phys. Rev.* **D74**, 112002 (2006).
24. V.M. Abazov *et al.* [DØ Collab.], *Phys. Rev. Lett.* **101**, 241801 (2008).
25. R. Aaij *et al.* [LHCb Collab.], *Eur. Phys. J.* **C72**, 2022 (2012).
26. R. Aaij *et al.* [LHCb Collab.], [arXiv:1602.07252 \[hep-ex\]](https://arxiv.org/abs/1602.07252).
27. R. Aaij *et al.* [LHCb Collab.], *Phys. Rev. Lett.* **114**, 041801 (2015).
28. R. Aaij *et al.* [LHCb Collab.], *Phys. Rev. Lett.* **113**, 211801 (2014).
29. B. Aubert *et al.* [BaBar Collab.], *Phys. Rev. Lett.* **94**, 161803 (2005); K.-F. Chen *et al.* [Belle Collab.], *Phys. Rev.* **D72**, 012004 (2005).
30. Throughout this document we omit references of results that have been replaced by new published measurements.
31. D. Buskulic *et al.* [ALEPH Collab.], *Z. Phys.* **C75**, 397 (1997).
32. P. Abreu *et al.* [DELPHI Collab.], *Z. Phys.* **C76**, 579 (1997).
33. M. Acciarri *et al.* [L3 Collab.], *Eur. Phys. J.* **C5**, 195 (1998).
34. G. Alexander *et al.* [OPAL Collab.], *Z. Phys.* **C72**, 377 (1996); K. Ackerstaff *et al.* [OPAL Collab.], *Z. Phys.* **C76**, 417 (1997); G. Abbiendi *et al.* [OPAL Collab.], *Phys. Lett.* **B493**, 266 (2000).
35. K. Ackerstaff *et al.* [OPAL Collab.], *Z. Phys.* **C76**, 401 (1997).
36. B. Aubert *et al.* [BaBar Collab.], *Phys. Rev. Lett.* **88**, 221802 (2002) and *Phys. Rev.* **D66**, 032003 (2002); *Phys. Rev. Lett.* **88**, 221803 (2002); *Phys. Rev.* **D67**, 072002 (2003); *Phys. Rev.* **D73**, 012004 (2006).
37. N.C. Hastings *et al.* [Belle Collab.], *Phys. Rev.* **D67**, 052004 (2003); Y. Zheng *et al.* [Belle Collab.], *Phys. Rev.* **D67**, 092004 (2003); K. Abe *et al.* [Belle Collab.], *Phys. Rev.* **D71**, 072003 (2005).
38. R. Aaij *et al.* [LHCb Collab.], *Phys. Lett.* **B709**, 177 (2012).
39. R. Aaij *et al.* [LHCb Collab.], *Phys. Lett.* **B719**, 318 (2013).
40. R. Aaij *et al.* [LHCb Collab.], *Eur. Phys. J.* **C73**, 2655 (2013).
41. R. Aaij *et al.* [LHCb Collab.], [arXiv:1604.03475 \[hep-ex\]](https://arxiv.org/abs/1604.03475).
42. H. Albrecht *et al.* [ARGUS Collab.], *Z. Phys.* **C55**, 357 (1992); *Phys. Lett.* **B324**, 249 (1994).

## Meson Particle Listings

 $B^0$ 

43. J. Bartelt *et al.* [CLEO Collab.], Phys. Rev. Lett. **71**, 1680 (1993).
44. B.H. Behrens *et al.* [CLEO Collab.], Phys. Lett. **B490**, 36 (2000).
45. B. Aubert *et al.* [BaBar Collab.], Phys. Rev. Lett. **92**, 181801 (2004) and Phys. Rev. **D70**, 012007 (2004).
46. T. Higuchi *et al.* [Belle Collab.], Phys. Rev. **D85**, 071105 (2012).
47. R. Aaij *et al.* [LHCb Collab.], JHEP **04**, 114 (2014).
48.  $\Delta m_d$  and  $\phi_s$  averages based only on data published before the end of February 2016 (*i.e.* without [41] for  $\Delta m_d$  and [66,67] for  $\phi_s$ ) can be found in the full listings of this Review.
49. See the review on the CKM quark-mixing matrix by A. Ceccucci, Z. Ligeti, and Y. Sakai in this publication.
50. S. Aoki *et al.* (FLAG working group) Eur. Phys. J. **C74**, 2890 (2014); see also <http://itpwiki.unibe.ch/flag/>.
51. R. Aaij *et al.* [LHCb Collab.], New J. Phys. **15**, 053021 (2013).
52. M. Bona *et al.* [UTfit Collab.], [arXiv:hep-ph/0606167v2](https://arxiv.org/abs/hep-ph/0606167v2); updated results at <http://www.utfit.org/>.
53. J. Charles *et al.* (CKMfitter Group), Phys. Rev. **D91**, 073007 (2015); updated results at <http://ckmfitter.in2p3.fr/>.
54. K. Hartkorn and H.-G. Moser, Eur. Phys. J. **C8**, 381 (1999).
55. M. Acciarri *et al.* [L3 Collab.], Phys. Lett. **B438**, 417 (1998).
56. D. Buskulic *et al.* [ALEPH Collab.], Phys. Lett. **B377**, 205 (1996); K. Akerstaff *et al.* [OPAL Collab.], Phys. Lett. **B426**, 161 (1998); F. Abe *et al.* [CDF Collab.], Phys. Rev. **D59**, 032004 (1999); V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. Lett. **114**, 062001 (2015); T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **107**, 272001 (2011); R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **113**, 172001 (2014).
57. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **112**, 111802 (2014).
58. R. Aaij *et al.* [LHCb Collab.], Phys. Lett. **B707**, 349 (2012); Phys. Lett. **B736**, 446 (2014).
59. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. **D84**, 052012 (2011); V.M. Abazov *et al.* [DØ Collab.], [arXiv:1603.01302 \[hep-ex\]](https://arxiv.org/abs/1603.01302).
60. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. **D87**, 112010 (2013).
61. R. Aaij *et al.* [LHCb Collab.], Nucl. Phys. **B873**, 275 (2013).
62. R. Fleischer and R. Knegjens, Eur. Phys. J. **C71**, 1789 (2011).
63. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. Lett. **109**, 171802 (2012).
64. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D85**, 032006 (2012).
65. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. **D90**, 052007 (2014).
66. G. Aad *et al.* [ATLAS Collab.], [arXiv:1601.03297 \[hep-ex\]](https://arxiv.org/abs/1601.03297).
67. V. Khachatryan *et al.* [CMS Collab.], Phys. Lett. **B757**, 97 (2016).
68. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **108**, 241801 (2012).
69. ALEPH, DELPHI, L3, OPAL, and SLD Collabs.; Physics Reports **427**, 257 (2006); we use the  $\bar{\chi}$  average given in Eq. (5.39).
70. D. Acosta *et al.* [CDF Collab.], Phys. Rev. **D69**, 012002 (2004).
71. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D74**, 092001 (2006).
72. P. Abreu *et al.* [DELPHI Collab.], Phys. Lett. **B289**, 199 (1992); P.D. Acton *et al.* [OPAL Collab.], Phys. Lett. **B295**, 357 (1992); D. Buskulic *et al.* [ALEPH Collab.], Phys. Lett. **B361**, 221 (1995).
73. P. Abreu *et al.* [DELPHI Collab.], Z. Phys. **C68**, 375 (1995); R. Barate *et al.* [ALEPH Collab.], Eur. Phys. J. **C2**, 197 (1998).
74. D. Buskulic *et al.* [ALEPH Collab.], Phys. Lett. **B384**, 449 (1996); J. Abdallah *et al.* [DELPHI Collab.], Eur. Phys. J. **C44**, 299 (2005).
75. R. Barate *et al.* [ALEPH Collab.], Eur. Phys. J. **C5**, 205 (1998).
76. J. Abdallah *et al.* [DELPHI Collab.], Phys. Lett. **B576**, 29 (2003).
77. F. Abe *et al.* [CDF Collab.], Phys. Rev. **D60**, 092005 (1999).
78. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. **D77**, 072003 (2008); T. Affolder *et al.* [CDF Collab.], Phys. Rev. Lett. **84**, 1663 (2000); the measurement of  $f_{\text{baryon}}/f_d$  in the latter paper has been updated based on T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. **D79**, 032001 (2009).
79. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. Lett. **99**, 052001 (2007); V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. Lett. **101**, 232002 (2008); T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. **D80**, 072003 (2009).
80. R. Aaij *et al.* [LHCb Collab.], JHEP **04**, 001 (2013).
81. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. **D85**, 032008 (2012).
82. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. Lett. **115**, 262001 (2015).
83. D.E. Jaffe *et al.* [CLEO Collab.], Phys. Rev. Lett. **86**, 5000 (2001).
84. J.P. Lees *et al.* [BaBar Collab.], Phys. Rev. Lett. **114**, 081801 (2015).
85. F. Abe *et al.* [CDF Collab.], Phys. Rev. **D55**, 2546 (1997).
86. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D89**, 012002 (2014).
87. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D86**, 072009 (2012).
88. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. Lett. **110**, 011801 (2013).
89. R. Aaij *et al.* [LHCb Collab.], Phys. Lett. **B728**, 607 (2014).
90. R. Barate *et al.* [ALEPH Collab.], Eur. Phys. J. **C20**, 431 (2001).
91. J.P. Lees *et al.* [BaBar Collab.], Phys. Rev. Lett. **111**, 101802 (2013).
92. E. Nakano *et al.* [Belle Collab.], Phys. Rev. **D73**, 112002 (2006).
93. G. Abbiendi *et al.* [OPAL Collab.], Eur. Phys. J. **C12**, 609 (2000).

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# Meson Particle Listings

## $B^0$

94. M. Beneke, G. Buchalla, and I. Dunietz, Phys. Lett. **B393**, 132 (1997); I. Dunietz, Eur. Phys. J. **C7**, 197 (1999).
95. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **114**, 041601 (2015).
96. R. Aaij *et al.* [LHCb Collab.], Phys. Lett. **B736**, 186 (2014).
97. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. **D86**, 052006 (2012).

### $B^0\text{-}\bar{B}^0$ MIXING PARAMETERS

For a discussion of  $B^0\text{-}\bar{B}^0$  mixing see the note on “ $B^0\text{-}\bar{B}^0$  Mixing” in the  $B^0$  Particle Listings above.

$\chi_d$  is a measure of the time-integrated  $B^0\text{-}\bar{B}^0$  mixing probability that a produced  $B^0(\bar{B}^0)$  decays as a  $\bar{B}^0(B^0)$ . Mixing violates  $\Delta B \neq 2$  rule.

$$\chi_d = \frac{x_d^2}{2(1+x_d^2)}$$

$$x_d = \frac{\Delta m_{B^0}}{\Gamma_{B^0}} = (m_{B_H^0} - m_{B_L^0}) \tau_{B^0},$$

where  $H, L$  stand for heavy and light states of two  $B^0$   $CP$  eigenstates and

$$\tau_{B^0} = \frac{1}{0.5(\Gamma_{B_H^0} + \Gamma_{B_L^0})}.$$

### $\chi_d$

This  $B^0\text{-}\bar{B}^0$  mixing parameter is the probability (integrated over time) that a produced  $B^0$  (or  $\bar{B}^0$ ) decays as a  $\bar{B}^0$  (or  $B^0$ ), e.g. for inclusive lepton decays

$$\chi_d = \frac{\Gamma(B^0 \rightarrow \ell^- X \text{ (via } \bar{B}^0))}{\Gamma(B^0 \rightarrow \ell^\pm X)} = \frac{\Gamma(\bar{B}^0 \rightarrow \ell^+ X \text{ (via } B^0))}{\Gamma(\bar{B}^0 \rightarrow \ell^\pm X)}$$

Where experiments have measured the parameter  $r = \chi/(1-\chi)$ , we have converted to  $\chi$ . Mixing violates the  $\Delta B \neq 2$  rule.

Note that the measurement of  $\chi$  at energies higher than the  $\Upsilon(4S)$  have not separated  $\chi_d$  from  $\chi_s$  where the subscripts indicate  $B^0(\bar{B}^0)$  or  $B_s^0(\bar{B}_s^0)$ . They are listed in the  $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE section.

The experiments at  $\Upsilon(4S)$  make an assumption about the  $B^0\text{-}\bar{B}^0$  fraction and about the ratio of the  $B^\pm$  and  $B^0$  semileptonic branching ratios (usually that it equals one).

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements, includes  $\chi_d$  calculated from  $\Delta m_{B^0}$  and  $\tau_{B^0}$ .

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>0.1875 \pm 0.0017</math> OUR EVALUATION</b>				
<b><math>0.182 \pm 0.015</math> OUR AVERAGE</b>				
$0.198 \pm 0.013 \pm 0.014$		<sup>1</sup> BEHRENS	00B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$0.16 \pm 0.04 \pm 0.04$		<sup>2</sup> ALBRECHT	94 ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$0.149 \pm 0.023 \pm 0.022$		<sup>3</sup> BARTELT	93 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
$0.171 \pm 0.048$		<sup>4</sup> ALBRECHT	92L ARG	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.20 \pm 0.13 \pm 0.12$		<sup>5</sup> ALBRECHT	96D ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$0.19 \pm 0.07 \pm 0.09$		<sup>6</sup> ALBRECHT	96D ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$0.24 \pm 0.12$		<sup>7</sup> ELSER	90 JADE	$e^+e^-$ 35–44 GeV
$0.158^{+0.052}_{-0.059}$		ARTUSO	89 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
$0.17 \pm 0.05$		<sup>8</sup> ALBRECHT	87I ARG	$e^+e^- \rightarrow \Upsilon(4S)$
$<0.19$	90	<sup>9</sup> BEAN	87B CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
$<0.27$	90	<sup>10</sup> AVERY	84 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> BEHRENS 00B uses high-momentum lepton tags and partially reconstructed  $\bar{B}^0 \rightarrow D^{*+}\pi^-, \rho^-$  decays to determine the flavor of the  $B$  meson.

<sup>2</sup> ALBRECHT 94 reports  $r=0.194 \pm 0.062 \pm 0.054$ . We convert to  $\chi$  for comparison. Uses tagged events (lepton + pion from  $D^*$ ).

<sup>3</sup> BARTELT 93 analysis performed using tagged events (lepton+pion from  $D^*$ ). Using dilepton events they obtain  $0.157 \pm 0.016^{+0.033}_{-0.028}$ .

<sup>4</sup> ALBRECHT 92L is a combined measurement employing several lepton-based techniques. It uses all previous ARGUS data in addition to new data and therefore supersedes ALBRECHT 87I. A value of  $r = 20.6 \pm 7.0\%$  is directly measured. The value can be used to measure  $x = \Delta M/\Gamma = 0.72 \pm 0.15$  for the  $B_d^0$  meson. Assumes  $\tau_{B^+}/\tau_{B^0} = 1.0 \pm 0.05$  and uses  $\tau_{B^\pm}/\tau_{B^0} = (0.95 \pm 0.14)(\tau_{B^\pm}/\tau_{B^0})$ .

<sup>5</sup> Uses  $D^{*+}K^\pm$  correlations.

<sup>6</sup> Uses  $(D^{*+}\ell^-)K^\pm$  correlations.

<sup>7</sup> These experiments see a combination of  $B_s$  and  $B_d$  mesons.

<sup>8</sup> ALBRECHT 87I is inclusive measurement with like-sign dileptons, with tagged  $B$  decays plus leptons, and one fully reconstructed event. Measures  $r=0.21 \pm 0.08$ . We convert to  $\chi$  for comparison. Superseded by ALBRECHT 92L.

<sup>9</sup> BEAN 87B measured  $r < 0.24$ ; we converted to  $\chi$ .

<sup>10</sup> Same-sign dilepton events. Limit assumes semileptonic BR for  $B^+$  and  $B^0$  equal. If  $B^0/B^\pm$  ratio  $< 0.58$ , no limit exists. The limit was corrected in BEAN 87B from  $r < 0.30$  to  $r < 0.37$ . We converted this limit to  $\chi$ .

$$\Delta m_{B^0} = m_{B_H^0} - m_{B_L^0}$$

$\Delta m_{B^0}$  is a measure of  $2\pi$  times the  $B^0\text{-}\bar{B}^0$  oscillation frequency in time-dependent mixing experiments.

The second “OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

The first “OUR EVALUATION”, also provided by the HFAG, includes  $\Delta m_d$  calculated from  $\chi_d$  measured at  $\Upsilon(4S)$ .

VALUE ( $10^{12} \text{ h s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.5096 \pm 0.0034</math> OUR EVALUATION</b>	First		
<b><math>0.5098 \pm 0.0035</math> OUR EVALUATION</b>	Second		
$0.503 \pm 0.011 \pm 0.013$	<sup>1</sup> AAIJ	13CF LHCb	$p\bar{p}$ at 7 TeV
$0.5156 \pm 0.0051 \pm 0.0033$	<sup>1</sup> AAIJ	13F LHCb	$p\bar{p}$ at 7 TeV
$0.499 \pm 0.032 \pm 0.003$	<sup>2</sup> AAIJ	12I LHCb	$p\bar{p}$ at 7 TeV
$0.506 \pm 0.020 \pm 0.016$	<sup>3</sup> ABZOV	06W D0	$p\bar{p}$ at 1.96 TeV
$0.511 \pm 0.007^{+0.007}_{-0.006}$	<sup>4</sup> AUBERT	06G BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.511 \pm 0.005 \pm 0.006$	<sup>5</sup> ABE	05B BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.531 \pm 0.025 \pm 0.007$	<sup>6</sup> ABDALLAH	03b DLPH	$e^+e^- \rightarrow Z$
$0.492 \pm 0.018 \pm 0.013$	<sup>7</sup> AUBERT	03C BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.503 \pm 0.008 \pm 0.010$	<sup>8</sup> HASTINGS	03 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.509 \pm 0.017 \pm 0.020$	<sup>9</sup> ZHENG	03 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.516 \pm 0.016 \pm 0.010$	<sup>10</sup> AUBERT	02I BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.493 \pm 0.012 \pm 0.009$	<sup>11</sup> AUBERT	02J BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.497 \pm 0.024 \pm 0.025$	<sup>12</sup> ABBIENDI, G	00b OPAL	$e^+e^- \rightarrow Z$
$0.503 \pm 0.064 \pm 0.071$	<sup>13</sup> ABE	99K CDF	$p\bar{p}$ at 1.8 TeV
$0.500 \pm 0.052 \pm 0.043$	<sup>14</sup> ABE	99Q CDF	$p\bar{p}$ at 1.8 TeV
$0.516 \pm 0.099^{+0.029}_{-0.035}$	<sup>15</sup> AFFOLDER	99c CDF	$p\bar{p}$ at 1.8 TeV
$0.471^{+0.078}_{-0.068}^{+0.033}_{-0.034}$	<sup>16</sup> ABE	98c CDF	$p\bar{p}$ at 1.8 TeV
$0.458 \pm 0.046 \pm 0.032$	<sup>17</sup> ACCIARRI	98D L3	$e^+e^- \rightarrow Z$
$0.437 \pm 0.043 \pm 0.044$	<sup>18</sup> ACCIARRI	98D L3	$e^+e^- \rightarrow Z$
$0.472 \pm 0.049 \pm 0.053$	<sup>19</sup> ACCIARRI	98D L3	$e^+e^- \rightarrow Z$
$0.523 \pm 0.072 \pm 0.043$	<sup>20</sup> ABREU	97N DLPH	$e^+e^- \rightarrow Z$
$0.493 \pm 0.042 \pm 0.027$	<sup>18</sup> ABREU	97N DLPH	$e^+e^- \rightarrow Z$
$0.499 \pm 0.053 \pm 0.015$	<sup>21</sup> ABREU	97N DLPH	$e^+e^- \rightarrow Z$
$0.480 \pm 0.040 \pm 0.051$	<sup>17</sup> ABREU	97N DLPH	$e^+e^- \rightarrow Z$
$0.444 \pm 0.029^{+0.020}_{-0.017}$	<sup>18</sup> ACKERSTAFF	97U OPAL	$e^+e^- \rightarrow Z$
$0.430 \pm 0.043^{+0.028}_{-0.030}$	<sup>17</sup> ACKERSTAFF	97V OPAL	$e^+e^- \rightarrow Z$
$0.482 \pm 0.044 \pm 0.024$	<sup>22</sup> BUSKULIC	97D ALEP	$e^+e^- \rightarrow Z$
$0.404 \pm 0.045 \pm 0.027$	<sup>18</sup> BUSKULIC	97D ALEP	$e^+e^- \rightarrow Z$
$0.452 \pm 0.039 \pm 0.044$	<sup>17</sup> BUSKULIC	97D ALEP	$e^+e^- \rightarrow Z$
$0.539 \pm 0.060 \pm 0.024$	<sup>23</sup> ALEXANDER	96V OPAL	$e^+e^- \rightarrow Z$
$0.567 \pm 0.089^{+0.029}_{-0.023}$	<sup>24</sup> ALEXANDER	96V OPAL	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.516 \pm 0.016 \pm 0.010$	<sup>25</sup> AUBERT	02N BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.494 \pm 0.012 \pm 0.015$	<sup>26</sup> HARA	02 BELL	Repl. by ABE 05B
$0.528 \pm 0.017 \pm 0.011$	<sup>27</sup> TOMURA	02 BELL	Repl. by ABE 05B
$0.463 \pm 0.008 \pm 0.016$	<sup>11</sup> ABE	01D BELL	Repl. by HASTINGS 03
$0.444 \pm 0.028 \pm 0.028$	<sup>28</sup> ACCIARRI	98D L3	$e^+e^- \rightarrow Z$
$0.497 \pm 0.035$	<sup>29</sup> ABREU	97N DLPH	$e^+e^- \rightarrow Z$
$0.467 \pm 0.022^{+0.017}_{-0.015}$	<sup>30</sup> ACKERSTAFF	97V OPAL	$e^+e^- \rightarrow Z$
$0.446 \pm 0.032$	<sup>31</sup> BUSKULIC	97D ALEP	$e^+e^- \rightarrow Z$
$0.531^{+0.050}_{-0.046} \pm 0.078$	<sup>32</sup> ABREU	96Q DLPH	Sup. by ABREU 97N
$0.496^{+0.055}_{-0.051} \pm 0.043$	<sup>17</sup> ACCIARRI	96E L3	Repl. by ACCIARRI 98D
$0.548 \pm 0.050^{+0.023}_{-0.019}$	<sup>33</sup> ALEXANDER	96V OPAL	$e^+e^- \rightarrow Z$
$0.496 \pm 0.046$	<sup>34</sup> AKERS	95J OPAL	Repl. by ACKERSTAFF 97V
$0.462^{+0.040}_{-0.053}^{+0.052}_{-0.035}$	<sup>17</sup> AKERS	95J OPAL	Repl. by ACKERSTAFF 97V
$0.50 \pm 0.12 \pm 0.06$	<sup>20</sup> ABREU	94M DLPH	Sup. by ABREU 97N
$0.508 \pm 0.075 \pm 0.025$	<sup>23</sup> AKERS	94C OPAL	Repl. by ALEXANDER 96V
$0.57 \pm 0.11 \pm 0.02$	<sup>24</sup> AKERS	94H OPAL	Repl. by ALEXANDER 96V
$0.50^{+0.07}_{-0.06}^{+0.11}_{-0.10}$	<sup>17</sup> BUSKULIC	94B ALEP	Sup. by BUSKULIC 97D
$0.52^{+0.10}_{-0.11}^{+0.04}_{-0.03}$	<sup>24</sup> BUSKULIC	93K ALEP	Sup. by BUSKULIC 97D

<sup>1</sup> Measured using  $B^0 \rightarrow D^-\pi^+$  and  $B^0 \rightarrow J/\psi K^*(892)^0$  decays.

<sup>2</sup> Measured using  $B^0 \rightarrow D^-\pi^+$ .

<sup>3</sup> Uses opposite-side flavor-tagging with  $B \rightarrow D^{(*)}\mu^+\mu^-X$  events.

<sup>4</sup> Measured using a simultaneous fit of the  $B^0$  lifetime and  $\bar{B}^0\text{-}B^0$  oscillation frequency  $\Delta m_d$  in the partially reconstructed  $B^0 \rightarrow D^{*-}\ell^+\nu$  decays.

<sup>5</sup> Measurement performed using a combined fit of  $CP$ -violation, mixing and lifetimes.

<sup>6</sup> Events with a high transverse momentum lepton were removed and an inclusively reconstructed vertex was required.

<sup>7</sup> AUBERT 03C uses a sample of approximately 14,000 exclusively reconstructed  $B^0 \rightarrow D^*(2010)^-\ell^+\nu$  and simultaneously measures the lifetime and oscillation frequency.

# Meson Particle Listings

## $B^0$

- <sup>8</sup> HASTINGS 03 measurement based on the time evolution of dilepton events. It also reports  $f_{+}/f_0 = 1.01 \pm 0.03 \pm 0.09$  and  $CPT$  violation parameters in  $B^0$ - $\bar{B}^0$  mixing.
- <sup>9</sup> ZHENG 03 data analyzed using partially reconstructed  $\bar{B}^0 \rightarrow D^{*-}\pi^+$  decay and a flavor tag based on the charge of the lepton from the accompanying  $B$  decay.
- <sup>10</sup> Uses a tagged sample of fully-reconstructed neutral  $B$  decays at  $\Upsilon(4S)$ .
- <sup>11</sup> Measured based on the time evolution of dilepton events in  $\Upsilon(4S)$  decays.
- <sup>12</sup> Data analyzed using partially reconstructed  $\bar{B}^0 \rightarrow D^{*+}\ell^-\bar{\nu}$  decay and a combination of flavor tags from the rest of the event.
- <sup>13</sup> Uses di-muon events.
- <sup>14</sup> Uses jet-charge and lepton-flavor tagging.
- <sup>15</sup> Uses  $\ell^+D^{*+}\ell^-$  events.
- <sup>16</sup> Uses  $\pi^-B$  in the same side.
- <sup>17</sup> Uses  $\ell\text{-}\ell$ .
- <sup>18</sup> Uses  $\ell\text{-}Q_{\text{hem}}$ .
- <sup>19</sup> Uses  $\ell\text{-}\ell$  with impact parameters.
- <sup>20</sup> Uses  $D^{*\pm}Q_{\text{hem}}$ .
- <sup>21</sup> Uses  $\pi_S^{\pm}\ell\text{-}Q_{\text{hem}}$ .
- <sup>22</sup> Uses  $D^{*\pm}\ell/Q_{\text{hem}}$ .
- <sup>23</sup> Uses  $D^{*\pm}\ell\text{-}Q_{\text{hem}}$ .
- <sup>24</sup> Uses  $D^{*\pm}\ell$ .
- <sup>25</sup> AUBERT 02N result based on the same analysis and data sample reported in AUBERT 02I.
- <sup>26</sup> Uses a tagged sample of  $B^0$  decays reconstructed in the mode  $B^0 \rightarrow D^{*}\ell\nu$ .
- <sup>27</sup> Uses a tagged sample of fully-reconstructed hadronic  $B^0$  decays at  $\Upsilon(4S)$ .
- <sup>28</sup> ACCIARRI 98D combines results from  $\ell\text{-}\ell$ ,  $\ell\text{-}Q_{\text{hem}}$ , and  $\ell\text{-}\ell$  with impact parameters.
- <sup>29</sup> ABREU 97N combines results from  $D^{*\pm}Q_{\text{hem}}$ ,  $\ell\text{-}Q_{\text{hem}}$ ,  $\pi_S^{\pm}\ell\text{-}Q_{\text{hem}}$ , and  $\ell\text{-}\ell$ .
- <sup>30</sup> ACKERSTAFF 97V combines results from  $\ell\text{-}\ell$ ,  $\ell\text{-}Q_{\text{hem}}$ ,  $D^{*}\ell$ , and  $D^{*\pm}Q_{\text{hem}}$ .
- <sup>31</sup> BUSKULIC 97D combines results from  $D^{*\pm}\ell/Q_{\text{hem}}$ ,  $\ell\text{-}Q_{\text{hem}}$ , and  $\ell\text{-}\ell$ .
- <sup>32</sup> ABREU 96Q analysis performed using lepton, kaon, and jet-charge tags.
- <sup>33</sup> ALEXANDER 96V combines results from  $D^{*\pm}\ell$  and  $D^{*\pm}\ell\text{-}Q_{\text{hem}}$ .
- <sup>34</sup> AKERS 95J combines results from charge measurement,  $D^{*\pm}\ell\text{-}Q_{\text{hem}}$  and  $\ell\text{-}\ell$ .

### $\chi_d = \Delta m_{B^0}/\Gamma_{B^0}$

The second “OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

The first “OUR EVALUATION”, also provided by the HFAG, includes  $\chi_d$  measured at  $\Upsilon(4S)$ .

VALUE	DOCUMENT ID
<b><math>0.775 \pm 0.006</math> OUR EVALUATION</b>	First
<b><math>0.775 \pm 0.006</math> OUR EVALUATION</b>	Second

### $\text{Re}(\chi_{CP} / |\lambda_{CP}|) \text{ Re}(z)$

The  $\chi_{CP}$  characterizes  $B^0$  and  $\bar{B}^0$  decays to states of charmonium plus  $K^0$ . Parameter  $z$  is used to describe  $CPT$  violation in mixing, see the review on “ $CP$  Violation” in the reviews section.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.014 \pm 0.035 \pm 0.034</math></b>	<sup>1</sup> AUBERT,B	04c	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Corresponds to 90% confidence range  $[-0.072, 0.101]$ .

### $\Delta\Gamma \text{ Re}(z)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.0071 \pm 0.0039 \pm 0.0020</math></b>	AUBERT	06T	BABR $e^+e^- \rightarrow \Upsilon(4S)$

### $\text{Re}(z)$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.9 \pm 3.7 \pm 3.3</math></b>	<sup>1</sup> HIGUCHI	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0 \pm 12 \pm 1$  <sup>2</sup> HASTINGS 03 BELL Repl. by HIGUCHI 12

<sup>1</sup> Measured using  $B^0 \rightarrow J/\psi K_S^0$ ,  $J/\psi K_L^0$ ,  $D^-\pi^+$ ,  $D^{*-}\pi^+$ ,  $D^{*-}\rho^+$ , and  $D^{*-}\ell^+\nu$  decays.

<sup>2</sup> Measured using inclusive dilepton events from  $B^0$  decay.

### $\text{Im}(z)$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>-0.8 \pm 0.4</math> OUR AVERAGE</b>			

$-0.57 \pm 0.33 \pm 0.33$  <sup>1</sup> HIGUCHI 12 BELL  $e^+e^- \rightarrow \Upsilon(4S)$

$-1.39 \pm 0.73 \pm 0.32$  <sup>2</sup> AUBERT 06T BABR  $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.8 \pm 2.9 \pm 2.5$  <sup>3</sup> AUBERT,B 04c BABR Repl. by AUBERT 06T

$-3 \pm 1 \pm 3$  <sup>4</sup> HASTINGS 03 BELL Repl. by HIGUCHI 12

<sup>1</sup> Measured using  $B^0 \rightarrow J/\psi K_S^0$ ,  $J/\psi K_L^0$ ,  $D^-\pi^+$ ,  $D^{*-}\pi^+$ ,  $D^{*-}\rho^+$ , and  $D^{*-}\ell^+\nu$  decays.

<sup>2</sup> Assuming  $\Delta\Gamma = 0$ , the result becomes  $\text{Im}(z) = -0.0037 \pm 0.0046$ .

<sup>3</sup> Corresponds to 90% confidence range  $[-0.028, 0.104]$ .

<sup>4</sup> Measured using inclusive dilepton events from  $B^0$  decay.

## CP VIOLATION PARAMETERS

### $\text{Re}(\epsilon_{B^0})/(1+|\epsilon_{B^0}|^2)$

$CP$  impurity in  $B_d^0$  system. It is obtained from either  $a_{\ell\ell}$ , the charge asymmetry in like-sign dilepton events or  $a_{CP}$ , the time-dependent asymmetry of inclusive  $B^0$  and  $\bar{B}^0$  decays.

The second “OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements. It assumes there is no  $CP$  violation in  $B_s$  mixing.

The first “OUR EVALUATION”, also provided by the HFAG, uses the measurements from  $B$ -factories only.

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>-0.4 \pm 0.4</math> OUR EVALUATION</b>	first eval		
<b><math>-0.2 \pm 0.5</math> OUR EVALUATION</b>	second eval		
<b><math>-0.1 \pm 0.4</math> OUR AVERAGE</b>			
$-0.05 \pm 0.48 \pm 0.75$	<sup>1</sup> AAIJ	15F	LHCB $p\bar{p}$ at 7, 8 TeV
$-0.975 \pm 0.875 \pm 0.475$	<sup>2</sup> LEES	15A	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$1.55 \pm 1.05$	<sup>3</sup> ABAZOV	14	D0 $p\bar{p}$ at 1.96 TeV
$0.15 \pm 0.42 \pm 0.94$	<sup>4</sup> LEES	13N	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-1.7 \pm 1.1 \pm 0.4$	<sup>5</sup> ABAZOV	12Ac	D0 $p\bar{p}$ at 1.96 TeV
$0.4 \pm 1.3 \pm 0.9$	<sup>6</sup> AUBERT	06T	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.3 \pm 2.0 \pm 2.1$	<sup>7</sup> NAKANO	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-3.2 \pm 6.5$	<sup>8</sup> BARATE	01D	ALEP $e^+e^- \rightarrow Z$
$3.5 \pm 10.3 \pm 1.5$	<sup>9</sup> JAFFE	01	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
$1.2 \pm 13.8 \pm 3.2$	<sup>10</sup> ABBIENDI	99J	OPAL $e^+e^- \rightarrow Z$
$2 \pm 7 \pm 3$	<sup>11</sup> ACKERSTAFF	97U	OPAL $e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.3 \pm 1.3$	<sup>12</sup> ABAZOV	11U	D0 Repl. by ABAZOV 14
$-2.3 \pm 1.1 \pm 0.8$	<sup>13</sup> ABAZOV	06s	D0 Repl. by ABAZOV 11U
$-14.7 \pm 6.7 \pm 5.7$	<sup>14</sup> AUBERT,B	04c	BABR Repl. by AUBERT 06T
$1.2 \pm 2.9 \pm 3.6$	<sup>2</sup> AUBERT	02k	BABR Repl. by LEES 15A
$4 \pm 18 \pm 3$	<sup>15</sup> BEHRENS	00B	CLE2 Repl. by JAFFE 01
$< 45$	<sup>16</sup> BARTELT	93	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 15F uses semileptonic  $B^0$  decays in the inclusive final states  $D^-\mu^+$  and  $D^{*-}\mu^+$ , where the  $D^-\mu^+$  meson decays into the  $K^+\pi^-\pi^+$  final state, and the  $D^{*-}$  meson into the  $\bar{D}^0(\rightarrow K^+\pi^-\pi^0)\pi^-$  final state. Reports  $A_{SL}^d = (-0.02 \pm 0.19 \pm 0.30)\%$ , which equals to  $4\text{Re}(\epsilon_{B^0})/(1+|\epsilon_{B^0}|^2)$ .

<sup>2</sup> Uses the charge asymmetry in like-sign dilepton events. LEES 15A reports  $A_{SL}^d = (-3.9 \pm 3.5 \pm 1.9) \times 10^{-3}$ .

<sup>3</sup> ABAZOV 14 uses the dimuon charge asymmetry with different impact parameters from which it reports  $A_{SL}^d = (-0.62 \pm 0.42) \times 10^{-2}$ .

<sup>4</sup> Uses  $B^0 \rightarrow D^{*-}X\ell^+\nu_\ell$  and a kaon-tagged sample which yields measurement of  $A_{SL}^d = (0.06 \pm 0.17 \pm 0.38 \pm 0.32)\%$ , corresponding to  $\Delta_{CP} = 1 - |q/p| = (0.29 \pm 0.84 \pm 1.88 \pm 1.61) \times 10^{-3}$ .

<sup>5</sup> ABAZOV 12Ac uses  $B^0 \rightarrow D^-\mu^+X$  and  $B^0 \rightarrow D^*(2010)^-\mu^+X$  decays without initial state flavor tagging which yields measurement of  $A_{SL}^d = (6.8 \pm 4.5 \pm 1.4) \times 10^{-3}$ .

<sup>6</sup> AUBERT 06T reports  $|q/p| - 1 = (-0.8 \pm 2.7 \pm 1.9) \times 10^{-3}$ . We convert to  $(1 - |q/p|^2)/4$ .

<sup>7</sup> Uses the charge asymmetry in like-sign dilepton events and reports  $|q/p| = 1.0005 \pm 0.0040 \pm 0.0043$ .

<sup>8</sup> BARATE 01D measured by investigating time-dependent asymmetries in semileptonic and fully inclusive  $B_d^0$  decays.

<sup>9</sup> JAFFE 01 finds  $a_{\ell\ell} = 0.013 \pm 0.050 \pm 0.005$  and combines with the previous BEHRENS 00B independent measurement.

<sup>10</sup> Data analyzed using the time-dependent asymmetry of inclusive  $B^0$  decay. The production flavor of  $B^0$  mesons is determined using both the jet charge and the charge of secondary vertex in the opposite hemisphere.

<sup>11</sup> ACKERSTAFF 97U assumes  $CPT$  and is based on measuring the charge asymmetry in a sample of  $B^0$  decays defined by lepton and  $Q_{\text{hem}}$  tags. If  $CPT$  is not invoked,  $\text{Re}(\epsilon_B) = -0.006 \pm 0.010 \pm 0.006$  is found. The indirect  $CPT$  violation parameter is determined to  $\text{Im}(\delta B) = -0.020 \pm 0.016 \pm 0.006$ .

<sup>12</sup> ABAZOV 11U uses the dimuon charge asymmetry with different impact parameters from which it reports  $A_{SL}^d = (-1.2 \pm 5.2) \times 10^{-3}$ .

<sup>13</sup> Uses the dimuon charge asymmetry.

<sup>14</sup> AUBERT 04c reports  $|q/p| = 1.029 \pm 0.013 \pm 0.011$  and we converted it to  $(1 - |q/p|^2)/4$ .

<sup>15</sup> BEHRENS 00B uses high-momentum lepton tags and partially reconstructed  $\bar{B}^0 \rightarrow D^{*+}\pi^-$ ,  $\rho^-$  decays to determine the flavor of the  $B$  meson.

<sup>16</sup> BARTELT 93 finds  $a_{\ell\ell} = 0.031 \pm 0.096 \pm 0.032$  which corresponds to  $|a_{\ell\ell}| < 0.18$ , which yields the above  $|\text{Re}(\epsilon_{B^0})/(1+|\epsilon_{B^0}|^2)|$ .

### $A_{T/CP}$

$A_{T/CP}$  is defined as

$$\frac{P(\bar{B}^0 \rightarrow B^0) - P(B^0 \rightarrow \bar{B}^0)}{P(\bar{B}^0 \rightarrow B^0) + P(B^0 \rightarrow \bar{B}^0)},$$

the  $CPT$  invariant asymmetry between the oscillation probabilities  $P(\bar{B}^0 \rightarrow B^0)$  and  $P(B^0 \rightarrow \bar{B}^0)$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.005 \pm 0.012 \pm 0.014</math></b>	<sup>1</sup> AUBERT	02k	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 02k uses the charge asymmetry in like-sign dilepton events.

See key on page 601

Meson Particle Listings  
 $B^0$  $A_{CP}(B^0 \rightarrow D^*(2010)^+ D^-)$  $A_{CP}$  is defined as

$$\frac{B(\bar{B}^0 \rightarrow \bar{T}) - B(B^0 \rightarrow f)}{B(\bar{B}^0 \rightarrow \bar{T}) + B(B^0 \rightarrow f)},$$

the CP-violation charge asymmetry of exclusive  $B^0$  and  $\bar{B}^0$  decay.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.037 \pm 0.034</math> OUR AVERAGE</b>			
$0.06 \pm 0.05 \pm 0.02$	ROHRKEN	12	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.008 \pm 0.048 \pm 0.013$	AUBERT	09C	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.07 \pm 0.08 \pm 0.04$	<sup>1</sup> AUSHEV	04	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.12 \pm 0.06 \pm 0.02$	AUBERT	07Ai	BABR Repl. by AUBERT 09C
$-0.03 \pm 0.10 \pm 0.02$	AUBERT,B	06A	BABR Repl. by AUBERT 07Ai
$-0.03 \pm 0.11 \pm 0.05$	AUBERT	03J	BABR Repl. by AUBERT,B 06B

<sup>1</sup> Combines results from fully and partially reconstructed  $B^0 \rightarrow D^{*\pm} D \pi$  decays. $A_{CP}(B^0 \rightarrow [K^+ K^-]_D K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.20 \pm 0.15 \pm 0.02</math></b>	AAIJ	14BN	LHCb $pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.45 \pm 0.23 \pm 0.02$	AAIJ	13L	LHCb Repl. by AAIJ 14BN

 $A_{CP}(B^0 \rightarrow [K^+ \pi^-]_D K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.03 \pm 0.04 \pm 0.02</math></b>	AAIJ	14BN	LHCb $pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.08 \pm 0.08 \pm 0.01$	AAIJ	13L	LHCb Repl. by AAIJ 14BN

 $R_d^+ = \Gamma(B^0 \rightarrow [\pi^+ K^-]_D K^{*0}) / \Gamma(B^0 \rightarrow [\pi^- K^+]_D K^{*0})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.03 \pm 0.01</math></b>	AAIJ	14BN	LHCb $pp$ at 7, 8 TeV

 $R_d^- = \Gamma(\bar{B}^0 \rightarrow [\pi^- K^+]_D K^{*0}) / \Gamma(\bar{B}^0 \rightarrow [\pi^+ K^-]_D K^{*0})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.03 \pm 0.01</math></b>	AAIJ	14BN	LHCb $pp$ at 7, 8 TeV

 $A_{CP}(B^0 \rightarrow [\pi^+ \pi^-]_D K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.09 \pm 0.22 \pm 0.02</math></b>	AAIJ	14BN	LHCb $pp$ at 7, 8 TeV

 $A_{CP}(B^0 \rightarrow K^+ \pi^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.082 \pm 0.006</math> OUR AVERAGE</b>			
$-0.083 \pm 0.013 \pm 0.004$	AALTONEN	14P	CDF $p\bar{p}$ at 1.96 TeV
$-0.080 \pm 0.007 \pm 0.003$	AAIJ	13Ax	LHCb $pp$ at 7 TeV
$-0.069 \pm 0.014 \pm 0.007$	DUH	13	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.107 \pm 0.016 \pm 0.006$	LEES	13D	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.086 \pm 0.023 \pm 0.009$	AALTONEN	11N	CDF $p\bar{p}$ at 1.96 TeV
$-0.04 \pm 0.16$	<sup>1</sup> CHEN	00	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.088 \pm 0.011 \pm 0.008$	AAIJ	12V	LHCb Repl. by AAIJ 13Ax
$-0.094 \pm 0.018 \pm 0.008$	LIN	08	BELL Repl. by DUH 13
$-0.107 \pm 0.018 \pm 0.007$	AUBERT	07Af	BABR Repl. by LEES 13D
$-0.013 \pm 0.078 \pm 0.012$	ABULENCIA,A	06D	CDF Repl. by AALTONEN 11N
$-0.088 \pm 0.035 \pm 0.013$	<sup>2</sup> CHAO	05A	BELL Repl. by CHAO 04B
$-0.133 \pm 0.030 \pm 0.009$	<sup>3</sup> AUBERT,B	04K	BABR Repl. by AUBERT 07Af
$-0.101 \pm 0.025 \pm 0.005$	<sup>4</sup> CHAO	04B	BELL Repl. by LIN 08
$-0.07 \pm 0.08 \pm 0.02$	<sup>5</sup> AUBERT	02D	BABR Repl. by AUBERT 02Q
$-0.102 \pm 0.050 \pm 0.016$	<sup>6</sup> AUBERT	02Q	BABR Repl. by AUBERT,B 04K
$-0.06 \pm 0.09 \pm 0.01$	<sup>7</sup> CASEY	02	BELL Repl. by CHAO 04B
$0.044 \pm 0.186 \pm 0.018$	<sup>8</sup> ABE	01K	BELL Repl. by CASEY 02
$-0.167 \pm 0.021$	<sup>9</sup> AUBERT	01E	BABR Repl. by AUBERT 02Q

<sup>1</sup> Corresponds to 90% confidence range  $-0.30 < A_{CP} < 0.22$ .<sup>2</sup> Corresponds to a 90% CL interval of  $-0.15 < A_{CP} < -0.03$ .<sup>3</sup> Based on a total signal yield of  $N(K^- \pi^+) + N(K^+ \pi^-) = 1606 \pm 51$  events.<sup>4</sup> CHAO 04B reports significance of 3.9 standard deviation for deviation of  $A_{CP}$  from zero.<sup>5</sup> Corresponds to 90% confidence range  $-0.21 < A_{CP} < 0.07$ .<sup>6</sup> Corresponds to 90% confidence range  $-0.188 < A_{CP} < -0.016$ .<sup>7</sup> Corresponds to 90% confidence range  $-0.21 < A_{CP} < +0.09$ .<sup>8</sup> Corresponds to 90% confidence range  $-0.25 < A_{CP} < 0.37$ .<sup>9</sup> Corresponds to 90% confidence range  $-0.35 < A_{CP} < -0.03$ . $A_{CP}(B^0 \rightarrow \eta' K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.07 \pm 0.18</math> OUR AVERAGE</b>			
$-0.22 \pm 0.29 \pm 0.07$	SATO	14	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.02 \pm 0.23 \pm 0.02$	DEL-AMO-SA...10A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.08 \pm 0.25 \pm 0.02$	<sup>1</sup> AUBERT	07E	BABR Repl. by DEL-AMO-SANCHEZ 10A

<sup>1</sup> Reports  $A_{CP}$  with the opposite sign convention. $A_{CP}(B^0 \rightarrow \eta' K_0^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.19 \pm 0.17 \pm 0.02</math></b>	DEL-AMO-SA...10A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \eta' K_2^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.14 \pm 0.18 \pm 0.02</math></b>	DEL-AMO-SA...10A	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \eta K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.19 \pm 0.05</math> OUR AVERAGE</b>			
$0.17 \pm 0.08 \pm 0.01$	WANG	07B	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.21 \pm 0.06 \pm 0.02$	AUBERT,B	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.02 \pm 0.11 \pm 0.02$	AUBERT,B	04D	BABR Repl. by AUBERT,B 06H

 $A_{CP}(B^0 \rightarrow \eta K_0^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.13 \pm 0.02</math></b>	AUBERT,B	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \eta K_2^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.07 \pm 0.19 \pm 0.02</math></b>	AUBERT,B	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow b_1 K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.07 \pm 0.12 \pm 0.02</math></b>	AUBERT	07Bi	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \omega K^{*0})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.45 \pm 0.25 \pm 0.02</math></b>	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \omega(K\pi)_0^{*0})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.07 \pm 0.09 \pm 0.02</math></b>	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \omega K_2^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.37 \pm 0.17 \pm 0.02</math></b>	AUBERT	09H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow K^+ \pi^- \pi^0)$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0 \pm 6</math> OUR AVERAGE</b>			
$-3.0^{+4.5}_{-5.1} \pm 5.5$	<sup>1</sup> AUBERT	08Aq	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$7 \pm 11 \pm 1$	<sup>2</sup> CHANG	04	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			
<sup>2</sup> Corresponds to 90% confidence range $-0.12 < A_{CP} < 0.26$ .			

 $A_{CP}(B^0 \rightarrow \rho^- K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.20 \pm 0.11</math> OUR AVERAGE</b>			
$0.20 \pm 0.09 \pm 0.08$	<sup>1</sup> LEES	11	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.22 \pm 0.22 \pm 0.06$	<sup>2</sup> CHANG	04	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-0.23 \pm 0.02$			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.11 \pm 0.14 \pm 0.07$	<sup>1</sup> AUBERT	08Aq	BABR Repl. by LEES 11
$-0.28 \pm 0.17 \pm 0.08$	<sup>3</sup> AUBERT	03T	BABR Repl. by AUBERT 08Aq

<sup>1</sup> Uses Dalitz plot analysis of  $B^0 \rightarrow K^+ \pi^- \pi^0$  decays.<sup>2</sup> Corresponds to 90% confidence range  $-0.18 < A_{CP} < 0.64$ .<sup>3</sup> The result reported corresponds to  $-A_{CP}$ . $A_{CP}(B^0 \rightarrow \rho(1450)^- K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.10 \pm 0.32 \pm 0.09</math></b>	<sup>1</sup> LEES	11	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			

 $A_{CP}(B^0 \rightarrow \rho(1700)^- K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.36 \pm 0.57 \pm 0.23</math></b>	<sup>1</sup> LEES	11	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			

 $A_{CP}(B^0 \rightarrow K^+ \pi^- \pi^0 \text{ nonresonant})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.10 \pm 0.16 \pm 0.08</math></b>	<sup>1</sup> LEES	11	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.23 \pm 0.19 \pm 0.11$	<sup>1</sup> AUBERT	08Aq	BABR Repl. by LEES 11

<sup>1</sup> Uses Dalitz plot analysis of  $B^0 \rightarrow K^+ \pi^- \pi^0$  decays. The quoted value is only for the flat part of the non-resonant component.

# Meson Particle Listings

$B^0$

$A_{CP}(B^0 \rightarrow K^0 \pi^+ \pi^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.01 \pm 0.05 \pm 0.01$	<sup>1</sup> AUBERT	09AU BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

$A_{CP}(B^0 \rightarrow K^*(892)^+ \pi^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.22 \pm 0.06$ OUR AVERAGE			
$-0.29 \pm 0.11 \pm 0.02$	<sup>1</sup> LEES	11 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.21 \pm 0.10 \pm 0.02$	<sup>2,3</sup> AUBERT	09AU BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.21 \pm 0.11 \pm 0.07$	<sup>4</sup> DALSENO	09 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.26 \pm^{+0.33}_{-0.34} \pm^{+0.10}_{-0.08}$	<sup>5</sup> EISENSTEIN	03 CLE2	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.19 \pm^{+0.20}_{-0.15} \pm 0.04$	<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
$-0.11 \pm 0.14 \pm 0.05$	<sup>2</sup> AUBERT	06I BABR	Repl. by AUBERT 09AU
$0.23 \pm 0.18 \pm^{+0.09}_{-0.06}$	AUBERT,B	04O BABR	Repl. by AUBERT 06I
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays.			
<sup>3</sup> The first of two equivalent solutions is used.			
<sup>4</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two consistent solutions that may be preferred.			
<sup>5</sup> Corresponds to 90% confidence range $-0.31 < A_{CP} < 0.78$ .			

$A_{CP}(B^0 \rightarrow (K\pi)_0^{*+} \pi^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.09 \pm 0.07$ OUR AVERAGE			
$0.07 \pm 0.14 \pm 0.01$	<sup>1</sup> LEES	11 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.09 \pm 0.07 \pm 0.03$	<sup>2</sup> AUBERT	09AU BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.17 \pm^{+0.11}_{-0.16} \pm 0.22$	<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

$A_{CP}(B^0 \rightarrow (K\pi)_0^{*0} \pi^0)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.15 \pm 0.10 \pm 0.04$	<sup>1</sup> LEES	11 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.22 \pm 0.12 \pm^{+0.30}_{-0.29}$	<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			

$A_{CP}(B^0 \rightarrow K^{*0} \pi^0)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.15 \pm 0.12 \pm 0.04$	<sup>1</sup> LEES	11 BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.09 \pm^{+0.21}_{-0.24} \pm 0.09$	<sup>1</sup> AUBERT	08AQ BABR	Repl. by LEES 11
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^+ \pi^- \pi^0$ decays.			

$A_{CP}(B^0 \rightarrow K^*(892)^0 \pi^+ \pi^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.07 \pm 0.04 \pm 0.03$	AUBERT	07As BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow K^*(892)^0 \rho^0)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.06 \pm 0.09 \pm 0.02$	LEES	12K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.09 \pm 0.19 \pm 0.02$	AUBERT,B	06G BABR	Repl. by LEES 12K

$A_{CP}(B^0 \rightarrow K^{*0} f_0(980))$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.07 \pm 0.10 \pm 0.02$	LEES	12K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.17 \pm 0.28 \pm 0.02$	AUBERT,B	06G BABR	Repl. by LEES 12K

$A_{CP}(B^0 \rightarrow K^{*+} \rho^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.21 \pm 0.15 \pm 0.02$	LEES	12K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow K^*(892)^0 K^+ K^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.01 \pm 0.05 \pm 0.02$	AUBERT	07As BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow a_1^- K^+)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.16 \pm 0.12 \pm 0.01$	AUBERT	08F BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow K^0 K^0)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.58 \pm^{+0.73}_{-0.66} \pm 0.04$	LIN	07 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow K^*(892)^0 \phi)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.00 \pm 0.04$ OUR AVERAGE			
$-0.007 \pm 0.048 \pm 0.021$	PRIM	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.01 \pm 0.06 \pm 0.03$	AUBERT	08Bg BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.03 \pm 0.07 \pm 0.03$	AUBERT	07D BABR	Repl. by AUBERT 08Bg
$0.02 \pm 0.09 \pm 0.02$	<sup>1</sup> CHEN	05A BELL	Repl. by PRIM 13
$-0.01 \pm 0.09 \pm 0.02$	AUBERT,B	04W BABR	Repl. by AUBERT 07D
$0.04 \pm 0.12 \pm 0.02$	AUBERT	03V BABR	Repl. by AUBERT 04W
$0.07 \pm 0.15 \pm^{+0.05}_{-0.03}$	<sup>2</sup> CHEN	03B BELL	Repl. by CHEN 05A
$0.00 \pm 0.27 \pm 0.03$	<sup>3</sup> AUBERT	02E BABR	Repl. by AUBERT 03V

<sup>1</sup> Corresponds to 90% confidence range  $-0.14 < A_{CP} < 0.17$ .  
<sup>2</sup> Corresponds to 90% confidence range  $-0.18 < A_{CP} < 0.33$ .  
<sup>3</sup> Corresponds to 90% confidence range  $-0.44 < A_{CP} < 0.44$ .

$A_{CP}(B^0 \rightarrow K^*(892)^0 K^- \pi^+)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.22 \pm 0.33 \pm 0.20$	AUBERT	07As BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow \phi(K\pi)_0^{*0})$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.12 \pm 0.08$ OUR AVERAGE			
$0.093 \pm 0.094 \pm 0.017$	PRIM	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.20 \pm 0.14 \pm 0.06$	AUBERT	08Bg BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.17 \pm 0.15 \pm 0.03$	AUBERT	07D BABR	Repl. by AUBERT 08Bg

$A_{CP}(B^0 \rightarrow \phi K_2^*(1430)^0)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.11 \pm 0.10$ OUR AVERAGE			
$-0.155 \pm^{+0.152}_{-0.133} \pm 0.033$	PRIM	13 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.08 \pm 0.12 \pm 0.05$	AUBERT	08Bg BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.12 \pm 0.14 \pm 0.04$	AUBERT	07D BABR	Repl. by AUBERT 08Bg

$A_{CP}(B^0 \rightarrow K^*(892)^0 \gamma)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.002 \pm 0.015$ OUR AVERAGE			
$0.008 \pm 0.017 \pm 0.009$	AAIJ	13 LHCB	$pp$ at 7 TeV
$-0.016 \pm 0.022 \pm 0.007$	AUBERT	09Ao BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow K_2^*(1430)^0 \gamma)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08 \pm 0.15 \pm 0.01$	AUBERT,B	04U BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

$A_{CP}(B^0 \rightarrow \rho^+ \pi^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$0.13 \pm 0.06$ OUR AVERAGE	Error includes scale factor of 1.1.		
$0.09 \pm^{+0.05}_{-0.06} \pm 0.04$	<sup>1</sup> LEES	13J BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.21 \pm 0.08 \pm 0.04$	<sup>1</sup> KUSAKA	07 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.03 \pm 0.07 \pm 0.04$	AUBERT	07AA BABR	Repl. by LEES 13J
$-0.02 \pm 0.16 \pm^{+0.05}_{-0.02}$	WANG	05 BELL	Repl. by KUSAKA 07
$-0.18 \pm 0.08 \pm 0.03$	AUBERT	03T BABR	Repl. by AUBERT 07AA

<sup>1</sup> Uses time-dependent Dalitz plot analysis of  $B^0 \rightarrow \pi^+ \pi^- \pi^0$  decays.

$A_{CP}(B^0 \rightarrow \rho^- \pi^+)$			
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08 \pm 0.08$ OUR AVERAGE			
$-0.12 \pm 0.08 \pm^{+0.04}_{-0.05}$	<sup>1</sup> LEES	13J BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.08 \pm 0.16 \pm 0.11$	<sup>1</sup> KUSAKA	07 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.37 \pm 0.16 \pm^{+0.09}_{-0.10}$	AUBERT	07AA BABR	Repl. by LEES 13J
$-0.53 \pm 0.29 \pm^{+0.09}_{-0.04}$	WANG	05 BELL	Repl. by KUSAKA 07

<sup>1</sup> Uses time-dependent Dalitz plot analysis of  $B^0 \rightarrow \pi^+ \pi^- \pi^0$  decays.

$A_{CP}(B^0 \rightarrow a_1(1260)^\pm \pi^\mp)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.07 \pm 0.06</math> OUR AVERAGE</b>			
$-0.06 \pm 0.05 \pm 0.07$	DALSENO 12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.07 \pm 0.07 \pm 0.02$	AUBERT 07o	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow b_1^- \pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.05 \pm 0.10 \pm 0.02</math></b>			
	AUBERT 07Bi	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \rho \bar{\rho} K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.05 \pm 0.12</math> OUR AVERAGE</b>			
$-0.08 \pm 0.20 \pm 0.02$	CHEN 08c	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.11 \pm 0.13 \pm 0.06$	AUBERT 07Av	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow \rho \bar{\Lambda} \pi^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.04 \pm 0.07</math> OUR AVERAGE</b>			
$0.10 \pm 0.10 \pm 0.02$	AUBERT 09Ac	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.02 \pm 0.10 \pm 0.03$	WANG 07c	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.05 \pm 0.10</math> OUR AVERAGE</b>			
$0.02 \pm 0.20 \pm 0.02$	AUBERT 09T	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.08 \pm 0.12 \pm 0.02$	WEI 09A	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow K^{*0} e^+ e^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.21 \pm 0.19 \pm 0.02</math></b>			
	WEI 09A	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

 $A_{CP}(B^0 \rightarrow K^{*0} \mu^+ \mu^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.034 \pm 0.024</math> OUR AVERAGE</b>			
$-0.035 \pm 0.024 \pm 0.003$	AAIJ 14AN	LHCB	$pp$ at 7, 8 TeV
$0.00 \pm 0.15 \pm 0.03$	WEI 09A	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.072 \pm 0.040 \pm 0.005$	AAIJ 13E	LHCB	Repl. by AAIJ 14AN

 $C_{D^*(2010)^- D^+}(B^0 \rightarrow D^*(2010)^- D^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.01 \pm 0.11</math> OUR AVERAGE</b>			
$-0.13 \pm 0.16 \pm 0.05$	<sup>1</sup> ROHRKEN 12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.00 \pm 0.17 \pm 0.03$	AUBERT 09c	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.23 \pm 0.25 \pm 0.06$	<sup>2</sup> AUSHEV 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.23 \pm 0.15 \pm 0.04$	AUBERT 07Ai	BABR	Repl. by AUBERT 09c
$0.17 \pm 0.24 \pm 0.04$	AUBERT,B 05Z	BABR	Repl. by AUBERT 07Ai
$-0.22 \pm 0.37 \pm 0.10$	AUBERT 03J	BABR	Repl. by AUBERT,B 05Z
<sup>1</sup> ROHRKEN 12 reports the measurements of $C = -0.01 \pm 0.11 \pm 0.04$ and $\Delta C = 0.12 \pm 0.11 \pm 0.03$ such that $C_{D^*(2010)^- D^+} = C - \Delta C$ .			
<sup>2</sup> Combines results from fully and partially reconstructed $B^0 \rightarrow D^{*\pm} D^\mp$ decays.			

 $S_{D^*(2010)^- D^+}(B^0 \rightarrow D^*(2010)^- D^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.72 \pm 0.15</math> OUR AVERAGE</b>			
$-0.65 \pm 0.22 \pm 0.07$	<sup>1</sup> ROHRKEN 12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.73 \pm 0.23 \pm 0.050$	AUBERT 09c	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.96 \pm 0.43 \pm 0.12$	<sup>2</sup> AUSHEV 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.44 \pm 0.22 \pm 0.06$	AUBERT 07Ai	BABR	Repl. by AUBERT 09c
$-0.29 \pm 0.33 \pm 0.07$	AUBERT,B 05Z	BABR	Repl. by AUBERT 07Ai
$-0.24 \pm 0.69 \pm 0.12$	AUBERT 03J	BABR	Repl. by AUBERT,B 05Z
<sup>1</sup> ROHRKEN 12 reports the measurements of $S = -0.78 \pm 0.15 \pm 0.05$ and $\Delta S = -0.13 \pm 0.15 \pm 0.04$ such that $S_{D^*(2010)^- D^+} = S - \Delta S$ .			
<sup>2</sup> Combines results from fully and partially reconstructed $B^0 \rightarrow D^{*\pm} D^\mp$ decays.			

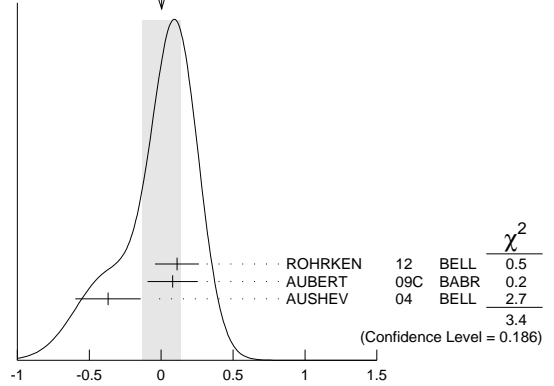
 $C_{D^*(2010)^+ D^-}(B^0 \rightarrow D^*(2010)^+ D^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.13</math> OUR AVERAGE</b>			Error includes scale factor of 1.3. See the ideogram below.
$0.11 \pm 0.14 \pm 0.06$	<sup>1</sup> ROHRKEN 12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.08 \pm 0.17 \pm 0.04$	AUBERT 09c	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.37 \pm 0.22 \pm 0.06$	<sup>2</sup> AUSHEV 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.18 \pm 0.15 \pm 0.04$	AUBERT 07Ai	BABR	Repl. by AUBERT 09c
$0.09 \pm 0.25 \pm 0.06$	AUBERT,B 05Z	BABR	Repl. by AUBERT 07Ai
$-0.47 \pm 0.40 \pm 0.12$	AUBERT 03J	BABR	Repl. by AUBERT,B 05Z

<sup>1</sup> ROHRKEN 12 reports the measurements of  $C = -0.01 \pm 0.11 \pm 0.04$  and  $\Delta C = 0.12 \pm 0.11 \pm 0.03$  such that  $C_{D^*(2010)^+ D^-} = C + \Delta C$ .

<sup>2</sup> Combines results from fully and partially reconstructed  $B^0 \rightarrow D^{*\pm} D^\mp$  decays.

WEIGHTED AVERAGE  
0.00±0.13 (Error scaled by 1.3)

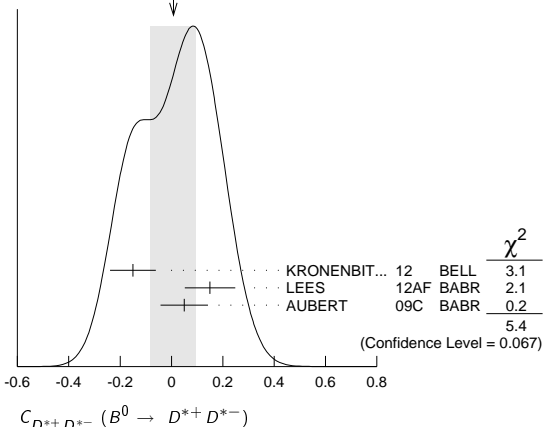
 $S_{D^*(2010)^+ D^-}(B^0 \rightarrow D^*(2010)^+ D^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.73 \pm 0.14</math> OUR AVERAGE</b>			
$-0.90 \pm 0.21 \pm 0.07$	<sup>1</sup> ROHRKEN 12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.62 \pm 0.21 \pm 0.03$	AUBERT 09c	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$-0.55 \pm 0.39 \pm 0.12$	<sup>2</sup> AUSHEV 04	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.79 \pm 0.21 \pm 0.06$	AUBERT 07Ai	BABR	Repl. by AUBERT 09c
$-0.54 \pm 0.35 \pm 0.07$	AUBERT,B 05Z	BABR	Repl. by AUBERT 07Ai
$-0.82 \pm 0.75 \pm 0.14$	AUBERT 03J	BABR	Repl. by AUBERT,B 05Z
<sup>1</sup> ROHRKEN 12 reports the measurements of $S = -0.78 \pm 0.15 \pm 0.05$ and $\Delta S = -0.13 \pm 0.15 \pm 0.04$ such that $S_{D^*(2010)^+ D^-} = S + \Delta S$ .			
<sup>2</sup> Combines results from fully and partially reconstructed $B^0 \rightarrow D^{*\pm} D^\mp$ decays.			

 $C_{D^{*+} D^{*-}}(B^0 \rightarrow D^{*+} D^{*-})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.09</math> OUR AVERAGE</b>			Error includes scale factor of 1.6. See the ideogram below.
$-0.15 \pm 0.08 \pm 0.04$	<sup>1,2</sup> KRONENBIT... 12	BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$+0.15 \pm 0.09 \pm 0.04$	<sup>3</sup> LEES 12AF	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.05 \pm 0.09 \pm 0.02$	AUBERT 09c	BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.15 \pm 0.13 \pm 0.04$	<sup>2</sup> VERVINK 09	BELL	Repl. by KRONENBITTER 12
$-0.02 \pm 0.11 \pm 0.02$	<sup>1</sup> AUBERT 07Bo	BABR	Repl. by AUBERT 09c
$0.26 \pm 0.26 \pm 0.06$	<sup>2</sup> MIYAKE 05	BELL	Repl. by VERVINK 09
$0.28 \pm 0.23 \pm 0.02$	<sup>4</sup> AUBERT 03Q	BABR	Repl. by AUBERT 07Bo
<sup>1</sup> Assumes both $CP$ -even and $CP$ -odd states having the $CP$ asymmetry.			
<sup>2</sup> Belle Collab. quotes $A_{D^{*+} D^{*-}}$ which is equal to $-C_{D^{*+} D^{*-}}$ .			
<sup>3</sup> Measured partially reconstructed candidates when one $D^0$ meson is not explicitly reconstructed. Analysis does not separate $CP$ -even and $CP$ -odd component.			
<sup>4</sup> AUBERT 03Q reports $ \lambda  = 0.75 \pm 0.19 \pm 0.02$ and $\text{Im}(\lambda) = 0.05 \pm 0.29 \pm 0.10$ . We convert them to $S$ and $C$ parameters taking into account correlations.			

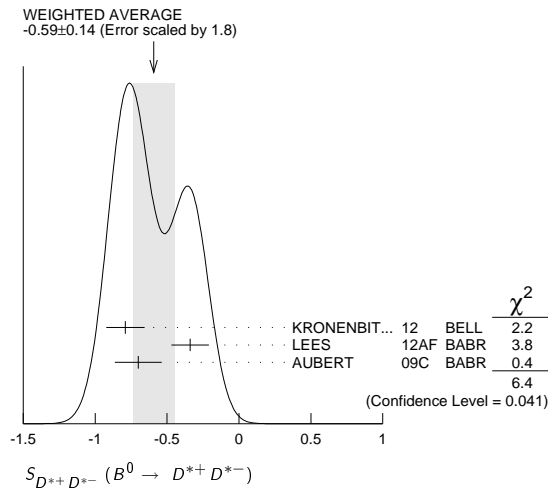
WEIGHTED AVERAGE  
0.01±0.09 (Error scaled by 1.6)



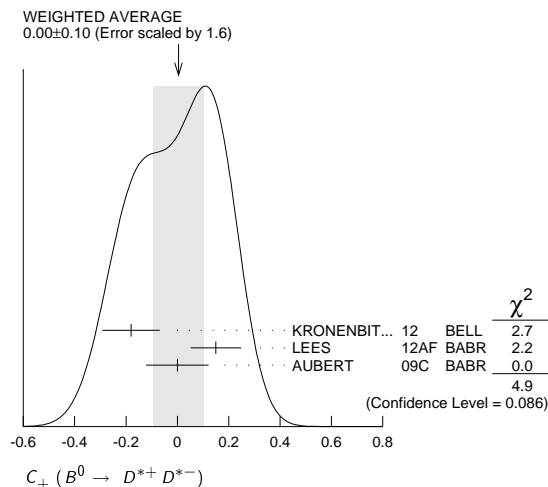
## Meson Particle Listings

 $B^0$  $S_{D^{*+}D^{*-}} (B^0 \rightarrow D^{*+}D^{*-})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.59 \pm 0.14</math> OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.		
$-0.79 \pm 0.13 \pm 0.03$	<sup>1</sup> KRONENBIT...12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.34 \pm 0.12 \pm 0.05$	<sup>2</sup> LEES	12AF BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.70 \pm 0.16 \pm 0.03$	<sup>1</sup> AUBERT	09c BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.96 \pm 0.25 \pm 0.13$	VERVINK	09 BELL	Repl. by KRONENBITTER 12
$-0.66 \pm 0.19 \pm 0.04$	<sup>1</sup> AUBERT	07Bo BABR	Repl. by AUBERT 09c
$-0.75 \pm 0.56 \pm 0.12$	MIYAKE	05 BELL	Repl. by VERVINK 09
$0.06 \pm 0.37 \pm 0.13$	<sup>3</sup> AUBERT	03Q BABR	Repl. by AUBERT 07Bo

<sup>1</sup> Assumes both  $CP$ -even and  $CP$ -odd states having the  $CP$  asymmetry.<sup>2</sup> Measured partially reconstructed candidates when one  $D^0$  meson is not explicitly reconstructed. Analysis does not separate  $CP$ -even and  $CP$ -odd component.<sup>3</sup> AUBERT 03Q reports  $|\lambda| = 0.75 \pm 0.19 \pm 0.02$  and  $\text{Im}(\lambda) = 0.05 \pm 0.29 \pm 0.10$ . We convert them to  $S$  and  $C$  parameters taking into account correlations. $C_+ (B^0 \rightarrow D^{*+}D^{*-})$ See the note in the  $C_{\pi\pi}$  datablock, but for  $CP$  even final state.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.10</math> OUR AVERAGE</b>	Error includes scale factor of 1.6. See the ideogram below.		
$-0.18 \pm 0.10 \pm 0.05$	<sup>1</sup> KRONENBIT...12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$+0.15 \pm 0.09 \pm 0.04$	<sup>2</sup> LEES	12AF BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.00 \pm 0.12 \pm 0.02$	AUBERT	09c BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.05 \pm 0.14 \pm 0.02$	AUBERT	07Bo BABR	Repl. by AUBERT 09c
$0.06 \pm 0.17 \pm 0.03$	<sup>3</sup> AUBERT, BE	05A BABR	Repl. by AUBERT 07Bo

<sup>1</sup> Belle Collab. quotes  $A_{D^{*+}D^{*-}}$  which is equal to  $-C_{D^{*+}D^{*-}}$ .<sup>2</sup> Measured partially reconstructed candidates when one  $D^0$  meson is not explicitly reconstructed. Extracted under assumption of equal  $C_+$  and  $C_-$ .<sup>3</sup> AUBERT, BE 05A reports a  $CP$ -odd fraction  $R_{\perp} = 0.125 \pm 0.044 \pm 0.007$ . $S_+ (B^0 \rightarrow D^{*+}D^{*-})$ See the note in the  $S_{\pi\pi}$  datablock, but for  $CP$  even final state.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.73 \pm 0.09</math> OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.		
$-0.81 \pm 0.13 \pm 0.03$	KRONENBIT...12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.49 \pm 0.18 \pm 0.08$	<sup>1</sup> LEES	12AF BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.76 \pm 0.16 \pm 0.04$	AUBERT	09c BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.72 \pm 0.19 \pm 0.05$	AUBERT	07Bo BABR	Repl. by AUBERT 09c
$-0.75 \pm 0.25 \pm 0.03$	<sup>2</sup> AUBERT, BE	05A BABR	Repl. by AUBERT 07Bo

<sup>1</sup> Measured partially reconstructed candidates when one  $D^0$  meson is not explicitly reconstructed. Analysis does not separate  $CP$ -even and  $CP$ -odd component. Value is obtained from  $S = -0.34 \pm 0.12 \pm 0.05$  using  $S = S_{\perp} (1 - 2R_{\perp})$  with  $R_{\perp} = 0.158 \pm 0.029$ .

<sup>2</sup> AUBERT, BE 05A reports a  $CP$ -odd fraction  $R_{\perp} = 0.125 \pm 0.044 \pm 0.007$ .

 $C_- (B^0 \rightarrow D^{*+}D^{*-})$ See the note in the  $C_{\pi\pi}$  datablock, but for  $CP$  odd final state.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.19 \pm 0.31</math> OUR AVERAGE</b>	Error includes scale factor of 3.5.		
$0.05 \pm 0.39 \pm 0.08$	<sup>1</sup> KRONENBIT...12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$0.41 \pm 0.49 \pm 0.08$	AUBERT	09c BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.23 \pm 0.67 \pm 0.10$	AUBERT	07Bo BABR	Repl. by AUBERT 09c
$-0.20 \pm 0.96 \pm 0.11$	<sup>2</sup> AUBERT, BE	05A BABR	Repl. by AUBERT 07Bo

<sup>1</sup> Belle Collab. quotes  $A_{D^{*+}D^{*-}}$  which is equal to  $-C_{D^{*+}D^{*-}}$ .

<sup>2</sup> AUBERT, BE 05A reports a  $CP$ -odd fraction  $R_{\perp} = 0.125 \pm 0.044 \pm 0.007$ .

 $S_- (B^0 \rightarrow D^{*+}D^{*-})$ See the note in the  $S_{\pi\pi}$  datablock, but for  $CP$  odd final state.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.1 \pm 1.6</math> OUR AVERAGE</b>	Error includes scale factor of 3.5.		
$1.52 \pm 0.62 \pm 0.12$	KRONENBIT...12	BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-1.80 \pm 0.70 \pm 0.16$	AUBERT	09c BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-1.83 \pm 1.04 \pm 0.23$	AUBERT	07Bo BABR	Repl. by AUBERT 09c
$-1.75 \pm 1.78 \pm 0.22$	<sup>1</sup> AUBERT, BE	05A BABR	Repl. by AUBERT 07Bo

<sup>1</sup> AUBERT, BE 05A reports a  $CP$ -odd fraction  $R_{\perp} = 0.125 \pm 0.044 \pm 0.007$ .

 $C (B^0 \rightarrow D^*(2010)^+ D^*(2010)^- K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.28 \pm 0.09</math></b>	<sup>1</sup> DALSENO	07 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

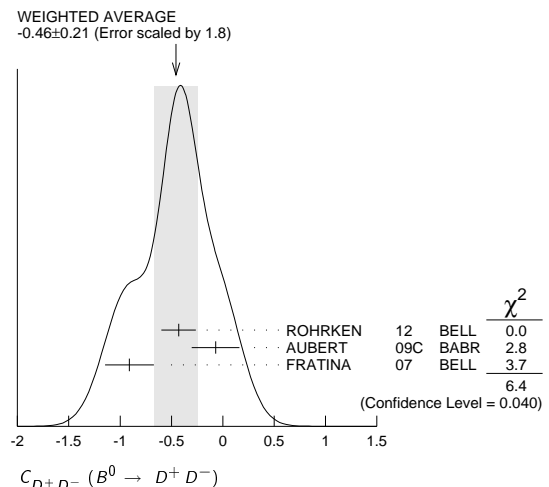
<sup>1</sup> Reports value of  $A$  which is equal to  $-C$ .

 $S (B^0 \rightarrow D^*(2010)^+ D^*(2010)^- K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.45 \pm 0.06</math></b>	<sup>1</sup> DALSENO	07 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> This value includes an unknown  $CP$  dilution factor  $D$  due to possible contributions from intermediate resonances and different partial waves. $C_{D^+D^-} (B^0 \rightarrow D^+D^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.46 \pm 0.21</math> OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.		
$-0.43 \pm 0.16 \pm 0.05$	ROHRKEN	12 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.07 \pm 0.23 \pm 0.03$	AUBERT	09c BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$-0.91 \pm 0.23 \pm 0.06$	<sup>1</sup> FRATINA	07 BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.11 \pm 0.22 \pm 0.07$	AUBERT	07A1 BABR	Repl. by AUBERT 09c
$0.11 \pm 0.35 \pm 0.06$	AUBERT, B	05Z BABR	Repl. by AUBERT 07A1

<sup>1</sup> The paper reports  $A$ , which is equal to  $-C$ .



$S_{D^+D^-}(B^0 \rightarrow D^+D^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.99 \pm 0.17</math> <math>-0.14</math> OUR AVERAGE</b>			
$-1.06 \pm 0.21$ $-0.14 \pm 0.08$	ROHRKEN	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.63 \pm 0.36 \pm 0.05$	AUBERT	09c	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-1.13 \pm 0.37 \pm 0.09$	FRATINA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.54 \pm 0.34 \pm 0.06$	AUBERT	07A1	BABR Repl. by AUBERT 09c
$-0.29 \pm 0.63 \pm 0.06$	AUBERT,B	05Z	BABR Repl. by AUBERT 07A1

 $C_{J/\psi(1S)\pi^0}(B^0 \rightarrow J/\psi(1S)\pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.13 \pm 0.13</math> OUR AVERAGE</b>			
$-0.20 \pm 0.19 \pm 0.03$	AUBERT	08AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.08 \pm 0.16 \pm 0.05$	<sup>1</sup> LEE	08A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.21 \pm 0.26 \pm 0.06$	AUBERT,B	06B	BABR Repl. by AUBERT 08AU
$0.01 \pm 0.29 \pm 0.03$	<sup>1</sup> KATAOKA	04	BELL Repl. by LEE 08A
$0.38 \pm 0.41 \pm 0.09$	AUBERT	03N	BABR Repl. by AUBERT,B 06B
<sup>1</sup> BELLE Collab. quotes $A_{J/\psi\pi^0}$ which is equal to $-C_{J/\psi\pi^0}$ .			

 $S_{J/\psi(1S)\pi^0}(B^0 \rightarrow J/\psi(1S)\pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.94 \pm 0.29</math> OUR AVERAGE</b>			Error includes scale factor of 1.9.
$-1.23 \pm 0.21 \pm 0.04$	AUBERT	08AU	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.65 \pm 0.21 \pm 0.05$	LEE	08A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.68 \pm 0.30 \pm 0.04$	AUBERT,B	06B	BABR Repl. by AUBERT 08AU
$-0.72 \pm 0.42 \pm 0.09$	KATAOKA	04	BELL Repl. by LEE 08A
$0.05 \pm 0.49 \pm 0.16$	AUBERT	03N	BABR Repl. by AUBERT,B 06B

 $C(B^0 \rightarrow J/\psi(1S)\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.063 \pm 0.056 \pm 0.019</math> <math>-0.014</math></b>	<sup>1</sup> AAIJ	15J	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Time-dependent $CP$ violation is measured in the $B^0 \rightarrow J/\psi\rho^0$ and was used to limit the size of penguin amplitude contributions to $\phi_S$ in $B_S^0 \rightarrow J/\psi\phi$ decays to be between $[-1.05^\circ, 1.18^\circ]$ at 95% confidence level.			

 $S(B^0 \rightarrow J/\psi(1S)\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.66 \pm 0.13 \pm 0.09</math> <math>-0.12 - 0.03</math></b>	<sup>1</sup> AAIJ	15J	LHCB $pp$ at 7, 8 TeV
<sup>1</sup> Time-dependent $CP$ violation is measured in the $B^0 \rightarrow J/\psi\rho^0$ and was used to limit the size of penguin amplitude contributions to $\phi_S$ in $B_S^0 \rightarrow J/\psi\phi$ decays to be between $[-1.05^\circ, 1.18^\circ]$ at 95% confidence level.			

 $C_{D_{CP}^{(*)}\rho^0}(B^0 \rightarrow D_{CP}^{(*)}\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.02 \pm 0.07 \pm 0.03</math></b>	<sup>1</sup> ABDESSALAM 15		$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.23 \pm 0.16 \pm 0.04$	AUBERT	07A1	BABR Repl. by ABDESSALAM 15
<sup>1</sup> BABAR and BELLE combined analysis uses $CP$ -eigenstate decay modes $D^0 \rightarrow K^+K^-$ , $K_S^0\pi^0$ , $K_S^0\omega$ , and $h^0 = \pi^0, \eta, \omega$ .			

 $S_{D_{CP}^{(*)}\rho^0}(B^0 \rightarrow D_{CP}^{(*)}\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.66 \pm 0.10 \pm 0.06</math></b>	<sup>1</sup> ABDESSALAM 15		$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.56 \pm 0.23 \pm 0.05$	AUBERT	07A1	BABR Repl. by ABDESSALAM 15
<sup>1</sup> BABAR and BELLE combined analysis uses $CP$ -eigenstate decay modes $D^0 \rightarrow K^+K^-$ , $K_S^0\pi^0$ , $K_S^0\omega$ , and $h^0 = \pi^0, \eta, \omega$ .			

 $C_{K^0\pi^0}(B^0 \rightarrow K^0\pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.13</math> OUR AVERAGE</b>			Error includes scale factor of 1.4.
$-0.14 \pm 0.13 \pm 0.06$	<sup>1</sup> FUJIKAWA	10A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.13 \pm 0.13 \pm 0.03$	AUBERT	09i	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.24 \pm 0.15 \pm 0.03$	AUBERT	08E	BABR Repl. by AUBERT 09i
$0.05 \pm 0.14 \pm 0.05$	<sup>1</sup> CHAO	07	BELL Repl. by FUJIKAWA 10A
$0.06 \pm 0.18 \pm 0.03$	AUBERT	05Y	BABR Repl. by AUBERT 08E
$-0.16 \pm 0.29 \pm 0.05$	<sup>1,2</sup> CHAO	05A	BELL Repl. by CHEN 05B
$0.11 \pm 0.20 \pm 0.09$	<sup>1</sup> CHEN	05B	BELL Repl. by CHAO 07
$-0.03 \pm 0.36 \pm 0.11$	<sup>1</sup> AUBERT	04M	BABR Repl. by AUBERT,B 04M
$0.40 \pm 0.27$ $-0.28 \pm 0.09$	<sup>3</sup> AUBERT,B	04M	BABR Repl. by AUBERT 05Y

<sup>1</sup> Reports  $A$  which is equal to  $-C$ .<sup>2</sup> Corresponds to a 90% CL interval of  $-0.33 < A_{CP} < 0.64$ .<sup>3</sup> Based on a total signal yield of  $122 \pm 16$  events. $S_{K^0\pi^0}(B^0 \rightarrow K^0\pi^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.58 \pm 0.17</math> OUR AVERAGE</b>			
$0.67 \pm 0.31 \pm 0.08$	FUJIKAWA	10A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.55 \pm 0.20 \pm 0.03$	AUBERT	09i	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.40 \pm 0.23 \pm 0.03$	AUBERT	08E	BABR Repl. by AUBERT 09i
$0.33 \pm 0.35 \pm 0.08$	CHAO	07	BELL Repl. by FUJIKAWA 10A
$0.35 \pm 0.30$ $-0.33 \pm 0.04$	AUBERT	05Y	BABR Repl. by AUBERT 08E
$0.32 \pm 0.61 \pm 0.13$	CHEN	05B	BELL Repl. by CHAO 07
$0.48 \pm 0.38$ $-0.47 \pm 0.06$	<sup>1</sup> AUBERT,B	04M	BABR Repl. by AUBERT 05Y

<sup>1</sup> Based on a total signal yield of  $122 \pm 16$  events. $C_{\eta'(958)K_S^0}(B^0 \rightarrow \eta'(958)K_S^0)$ See updated measurements in  $C_{\eta'K^0}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.04 \pm 0.20</math> OUR AVERAGE</b>			Error includes scale factor of 2.5.
$-0.21 \pm 0.10 \pm 0.02$	AUBERT	05M	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.19 \pm 0.11 \pm 0.05$	<sup>1</sup> CHEN	05B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.26 \pm 0.22 \pm 0.03$	<sup>1</sup> ABE	03C	BELL Repl. by ABE 03H
$0.01 \pm 0.16 \pm 0.04$	<sup>1</sup> ABE	03H	BELL Repl. by CHEN 05B
$0.10 \pm 0.22 \pm 0.04$	AUBERT	03W	BABR Repl. by AUBERT 05M
$-0.13 \pm 0.32 \pm 0.06$ $-0.09$	<sup>1</sup> CHEN	02B	BELL Repl. by ABE 03C

<sup>1</sup> BELLE Collab. quotes  $A_{\eta'(958)K_S^0}$  which is equal to  $-C_{\eta'(958)K_S^0}$ . $S_{\eta'(958)K_S^0}(B^0 \rightarrow \eta'(958)K_S^0)$ See updated measurements in  $S_{\eta'K^0}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.43 \pm 0.17</math> OUR AVERAGE</b>			Error includes scale factor of 1.5.
$0.30 \pm 0.14 \pm 0.02$	AUBERT	05M	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.65 \pm 0.18 \pm 0.04$	CHEN	05B	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.71 \pm 0.37 \pm 0.05$ $-0.06$	ABE	03C	BELL Repl. by ABE 03H
$0.43 \pm 0.27 \pm 0.05$	ABE	03H	BELL Repl. by CHEN 05B
$0.02 \pm 0.34 \pm 0.03$	AUBERT	03W	BABR Repl. by AUBERT 05M
$0.28 \pm 0.55 \pm 0.07$ $-0.08$	CHEN	02B	BELL Repl. by ABE 03C

 $C_{\eta'K^0}(B^0 \rightarrow \eta'K^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.06 \pm 0.04</math> OUR AVERAGE</b>			
$-0.03 \pm 0.05 \pm 0.04$	<sup>1</sup> SANTELJ	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.08 \pm 0.06 \pm 0.02$	AUBERT	09i	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.16 \pm 0.07 \pm 0.03$	<sup>2</sup> AUBERT	07A	BABR Repl. by AUBERT 09i
$0.01 \pm 0.07 \pm 0.05$	<sup>1,2</sup> CHEN	07	BELL Repl. by SANTELJ 14

<sup>1</sup> The paper reports  $A$ , which is equal to  $-C$ .<sup>2</sup> The mixing-induced  $CP$  violation is reported with a significance of more than 5 standard deviations in this  $b \rightarrow s$  penguin dominated mode. $S_{\eta'K^0}(B^0 \rightarrow \eta'K^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.63 \pm 0.06</math> OUR AVERAGE</b>			
$0.68 \pm 0.07 \pm 0.03$	SANTELJ	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.57 \pm 0.08 \pm 0.02$	AUBERT	09i	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.58 \pm 0.10 \pm 0.03$	<sup>1</sup> AUBERT	07A	BABR Repl. by AUBERT 09i
$0.64 \pm 0.10 \pm 0.04$	<sup>1</sup> CHEN	07	BELL Repl. by SANTELJ 14

<sup>1</sup> The mixing-induced  $CP$  violation is reported with a significance of more than 5 standard deviations in this  $b \rightarrow s$  penguin dominated mode. $C_{\omega K_S^0}(B^0 \rightarrow \omega K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0 \pm 0.4</math> OUR AVERAGE</b>			Error includes scale factor of 3.0.
$0.36 \pm 0.19 \pm 0.05$	<sup>1</sup> CHOBANOVA	14	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.52 \pm 0.22 \pm 0.03$ $-0.20$	AUBERT	09i	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.09 \pm 0.29 \pm 0.06$	<sup>1</sup> CHAO	07	BELL Repl. by CHOBANOVA 14
$-0.55 \pm 0.28 \pm 0.03$ $-0.26$	AUBERT,B	06E	BABR Repl. by AUBERT 09i
$-0.27 \pm 0.48 \pm 0.15$	<sup>1</sup> CHEN	05B	BELL Repl. by CHAO 07

<sup>1</sup> Belle Collab. quotes  $A_{\omega K_S^0}$  which is equal to  $-C_{\omega K_S^0}$ .

# Meson Particle Listings

$B^0$

## $S_{\omega K_S^0} (B^0 \rightarrow \omega K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.70 \pm 0.21</math> OUR AVERAGE</b>			
$0.91 \pm 0.32 \pm 0.05$	CHOBANOVA 14	BELL	$e^+ e^- \rightarrow \gamma(4S)$
$0.55^{+0.26}_{-0.29} \pm 0.02$	AUBERT 09i	BABR	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.11 \pm 0.46 \pm 0.07$	CHAO 07	BELL	Repl. by CHOBANOVA 14
$0.51^{+0.35}_{-0.39} \pm 0.02$	AUBERT,B 06e	BABR	Repl. by AUBERT 09i
$0.76 \pm 0.65^{+0.13}_{-0.16}$	CHEN 05b	BELL	Repl. by CHAO 07

## $C (B^0 \rightarrow K_S^0 \pi^0 \pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.23 \pm 0.52 \pm 0.13</math></b>	AUBERT	07AQ	BABR $e^+ e^- \rightarrow \gamma(4S)$

## $S (B^0 \rightarrow K_S^0 \pi^0 \pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.72 \pm 0.71 \pm 0.08</math></b>	AUBERT	07AQ	BABR $e^+ e^- \rightarrow \gamma(4S)$

## $C_{\rho K_S^0} (B^0 \rightarrow \rho^0 K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.04 \pm 0.20</math> OUR AVERAGE</b>			
$-0.05 \pm 0.26 \pm 0.10$	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
$-0.03 \pm 0.24_{-0.23} \pm 0.15$	<sup>2,3</sup> DALSENO 09	BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.64 \pm 0.41 \pm 0.20$	AUBERT 07F	BABR	Repl. by AUBERT 09AU
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			
<sup>2</sup> Quotes $A_{\rho^0(KS)^0}$ which is equal to $-C_{\rho^0 K_S^0}$ .			
<sup>3</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two consistent solutions that may be preferred.			

## $S_{\rho K_S^0} (B^0 \rightarrow \rho^0 K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.50^{+0.17}_{-0.21}</math> OUR AVERAGE</b>			
$0.35^{+0.26}_{-0.31} \pm 0.07$	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
$0.64^{+0.19}_{-0.25} \pm 0.13$	<sup>2</sup> DALSENO 09	BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.20 \pm 0.52 \pm 0.24$	AUBERT 07F	BABR	Repl. by AUBERT 09AU
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two consistent solutions that may be preferred.			

## $C_{f_0(980) K_S^0} (B^0 \rightarrow f_0(980) K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.29 \pm 0.20</math> OUR AVERAGE</b>			
$0.28 \pm 0.24 \pm 0.09$	<sup>1</sup> LEES 12o	BABR	$e^+ e^- \rightarrow \gamma(4S)$
$0.30 \pm 0.29 \pm 0.14$	<sup>2,3</sup> NAKAHAMA 10	BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.08 \pm 0.19 \pm 0.05$	<sup>4</sup> AUBERT 09AU	BABR	Repl. by LEES 12o
$0.06 \pm 0.17 \pm 0.11$	<sup>2,5</sup> DALSENO 09	BELL	Repl. by NAKAHAMA 10
$-0.41 \pm 0.23 \pm 0.07$	<sup>2</sup> AUBERT 07AX	BABR	Repl. by AUBERT 09AU
$0.15 \pm 0.15 \pm 0.07$	<sup>2</sup> CHAO 07	BELL	Repl. by DALSENO 09
$0.39 \pm 0.27 \pm 0.09$	<sup>2</sup> CHEN 05B	BELL	Repl. by CHAO 07
<sup>1</sup> Uses Dalitz plot analysis of the $B^0 \rightarrow K_S^0 K^+ K^-$ decay.			
<sup>2</sup> Quotes $A_{f_0(980) K_S^0}$ which is equal to $-C_{f_0(980) K_S^0}$ .			
<sup>3</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K_S^0 K^+ K^-$ decays and the first of four consistent solutions that may be preferred.			
<sup>4</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			
<sup>5</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two consistent solutions that may be preferred.			

## $S_{f_0(980) K_S^0} (B^0 \rightarrow f_0(980) K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.50 \pm 0.16</math> OUR AVERAGE</b>			
$-0.55 \pm 0.18 \pm 0.12$	<sup>1</sup> LEES 12o	BABR	$e^+ e^- \rightarrow \gamma(4S)$
$-0.43^{+0.22}_{-0.20} \pm 0.14$	<sup>2</sup> DALSENO 09	BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.96^{+0.21}_{-0.04} \pm 0.04$	<sup>3</sup> AUBERT 09AU	BABR	Repl. by LEES 12o
$-0.25 \pm 0.26 \pm 0.10$	<sup>4</sup> AUBERT 07AX	BABR	Repl. by AUBERT 09AU
$0.18 \pm 0.23 \pm 0.11$	CHAO 07	BELL	Repl. by DALSENO 09
$0.47 \pm 0.41 \pm 0.08$	CHEN 05B	BELL	Repl. by CHAO 07
<sup>1</sup> Uses Dalitz plot analysis of the $B^0 \rightarrow K_S^0 K^+ K^-$ decay.			
<sup>2</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two consistent solutions that may be preferred.			
<sup>3</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			
<sup>4</sup> Reports $\beta_{eff}$ . We quote $S$ obtained from epaps: E-PRLTAO-99-076741.			

## $S_{f_2(1270) K_S^0} (B^0 \rightarrow f_2(1270) K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.48 \pm 0.52 \pm 0.12</math></b>	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

## $C_{f_2(1270) K_S^0} (B^0 \rightarrow f_2(1270) K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.28^{+0.35}_{-0.40} \pm 0.11</math></b>	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

## $S_{f_x(1300) K_S^0} (B^0 \rightarrow f_x(1300) K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.20 \pm 0.52 \pm 0.10</math></b>	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

## $C_{f_x(1300) K_S^0} (B^0 \rightarrow f_x(1300) K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.13^{+0.33}_{-0.35} \pm 0.10</math></b>	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

## $S_{K^0 \pi^+ \pi^-} (B^0 \rightarrow K^0 \pi^+ \pi^- \text{ nonresonant})$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.01 \pm 0.31 \pm 0.10</math></b>	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

## $C_{K^0 \pi^+ \pi^-} (B^0 \rightarrow K^0 \pi^+ \pi^- \text{ nonresonant})$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.25 \pm 0.08</math></b>	<sup>1</sup> AUBERT 09AU	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

## $C_{K_S^0 K_S^0} (B^0 \rightarrow K_S^0 K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.0 \pm 0.4</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.		
$0.38 \pm 0.38 \pm 0.05$	<sup>1</sup> NAKAHAMA 08	BELL	$e^+ e^- \rightarrow \gamma(4S)$
$-0.40 \pm 0.41 \pm 0.06$	AUBERT,BE 06c	BABR	$e^+ e^- \rightarrow \gamma(4S)$
<sup>1</sup> Reports $A_{K_S^0 K_S^0}$ which equals to $-C_{K_S^0 K_S^0}$ .			

## $S_{K_S^0 K_S^0} (B^0 \rightarrow K_S^0 K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.8 \pm 0.5</math> OUR AVERAGE</b>			
$-0.38^{+0.69}_{-0.77} \pm 0.09$	NAKAHAMA 08	BELL	$e^+ e^- \rightarrow \gamma(4S)$
$-1.28^{+0.80}_{-0.73} \pm 0.16$	AUBERT,BE 06c	BABR	$e^+ e^- \rightarrow \gamma(4S)$

## $C_{K^+ K^- K_S^0} (B^0 \rightarrow K^+ K^- K_S^0 \text{ nonresonant})$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.08</math> OUR AVERAGE</b>			
$0.02 \pm 0.09 \pm 0.03$	<sup>1,2</sup> LEES 12o	BABR	$e^+ e^- \rightarrow \gamma(4S)$
$0.14 \pm 0.11 \pm 0.09$	<sup>3,4</sup> NAKAHAMA 10	BELL	$e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.054 \pm 0.102 \pm 0.060$	<sup>3,5</sup> AUBERT 07AX	BABR	Repl. by LEES 12o
$0.09 \pm 0.10 \pm 0.05$	<sup>3,5</sup> CHAO 07	BELL	Repl. by NAKAHAMA 10
$0.10 \pm 0.14 \pm 0.04$	<sup>5</sup> AUBERT 05T	BABR	Repl. by AUBERT 07AX
$0.09 \pm 0.12 \pm 0.07$	<sup>3</sup> CHEN 05B	BELL	Repl. by CHAO 07
$-0.10 \pm 0.19 \pm 0.10$	<sup>5</sup> AUBERT,B 04V	BABR	Repl. by AUBERT 05T
$0.40 \pm 0.33^{+0.28}_{-0.10}$	<sup>3</sup> ABE 03C	BELL	Repl. by ABE 03H
$0.17 \pm 0.16 \pm 0.04$	<sup>3,5</sup> ABE 03H	BELL	Repl. by CHEN 05B
<sup>1</sup> Uses Dalitz plot analysis of the $B^0 \rightarrow K_S^0 K^+ K^-$ decay.			
<sup>2</sup> This measurement is performed on all the isobar components, excluding $\phi K_S^0$ and $f_0(980) K_S^0$ .			
<sup>3</sup> Quotes $A_{K^+ K^- K_S^0}$ which is equal to $-C_{K^+ K^- K_S^0}$ .			
<sup>4</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K_S^0 K^+ K^-$ decays and the first of four consistent solutions that may be preferred.			
<sup>5</sup> Excludes the events from $B^0 \rightarrow \phi K_S^0$ decay. The results are derived from a combined sample of $K^+ K^- K_S^0$ and $K^+ K^- K_L^0$ decays.			

$S_{K^+K^-K_S^0} (B^0 \rightarrow K^+K^-K_S^0 \text{ nonresonant})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.66 \pm 0.11</math> OUR AVERAGE</b>			
$-0.65 \pm 0.12 \pm 0.03$	<sup>1,2</sup> LEES	12o BABR	$e^+e^- \rightarrow \gamma(4S)$
$-0.68 \pm 0.15 \pm 0.21 \pm 0.13$	<sup>3</sup> CHAO	07 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.764 \pm 0.111 \pm 0.071 \pm 0.040$	<sup>3,4</sup> AUBERT	07AX BABR	Repl. by LEES 12o
$-0.42 \pm 0.17 \pm 0.03$	<sup>3,5</sup> AUBERT	05T BABR	Repl. by AUBERT 07AX
$-0.49 \pm 0.18 \pm 0.04$	CHEN	05B BELL	Repl. by CHAO 07
$-0.56 \pm 0.25 \pm 0.04$	<sup>3,6</sup> AUBERT,B	04V BABR	Repl. by AUBERT 05T
$-0.49 \pm 0.43 \pm 0.11$	ABE	03C BELL	Repl. by ABE 03H
$-0.51 \pm 0.26 \pm 0.05$	<sup>3,7</sup> ABE	03H BELL	Repl. by CHEN 05B

<sup>1</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0 K^+ K^-$  decay.

<sup>2</sup> This measurement is performed on all the isobar components, excluding  $\phi K_S^0$  and  $f_0(980) K_S^0$ . Note that the nonresonant component is not a  $CP$  eigenstate.

<sup>3</sup> Excludes events from  $B^0 \rightarrow \phi K_S^0$  decay. The results are derived from a combined sample of  $K^+ K^- K_S^0$  and  $K^+ K^- K_L^0$  decays.

<sup>4</sup> Reports  $\beta_{eff}$ . We quote  $S$  obtained from epaps: E-PRLTAO-99-076741.

<sup>5</sup> The measured  $CP$ -even final states fraction is  $0.89 \pm 0.08 \pm 0.06$ .

<sup>6</sup> The measured  $CP$ -even final states fraction is  $0.98 \pm 0.15 \pm 0.04$ .

<sup>7</sup> The measured  $CP$ -even final states fraction is  $1.03 \pm 0.15 \pm 0.05$ .

 $C_{K^+K^-K_S^0} (B^0 \rightarrow K^+K^-K_S^0 \text{ inclusive})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.015 \pm 0.077 \pm 0.053</math></b>	<sup>1,2</sup> AUBERT	07AX BABR	$e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Measured using full Dalitz plot fit including  $\phi$  component.

<sup>2</sup> The results are derived from a combined sample of  $K^+ K^- K_S^0$  and  $K^+ K^- K_L^0$  decays.

 $S_{K^+K^-K_S^0} (B^0 \rightarrow K^+K^-K_S^0 \text{ inclusive})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.647 \pm 0.116 \pm 0.040</math></b>	<sup>1</sup> AUBERT	07AX BABR	$e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Measured using full Dalitz plot fit including  $\phi$  component.

 $C_{\phi K_S^0} (B^0 \rightarrow \phi K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.14</math> OUR AVERAGE</b>			

$0.05 \pm 0.18 \pm 0.05$	<sup>1</sup> LEES	12o BABR	$e^+e^- \rightarrow \gamma(4S)$
$-0.04 \pm 0.20 \pm 0.10$	<sup>2,3</sup> NAKAHAMA	10 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.08 \pm 0.18 \pm 0.04$	<sup>2,4</sup> AUBERT	07AX BABR	Repl. by LEES 12o
$-0.07 \pm 0.15 \pm 0.05$	<sup>2,4</sup> CHEN	07 BELL	Repl. by NAKAHAMA 10
$0.00 \pm 0.23 \pm 0.05$	<sup>4</sup> AUBERT	05T BABR	Repl. by AUBERT 07AX
$-0.08 \pm 0.22 \pm 0.09$	<sup>2,4</sup> CHEN	05B BELL	Repl. by CHEN 07
$0.01 \pm 0.33 \pm 0.10$	<sup>4</sup> AUBERT,B	04G BABR	Repl. by AUBERT 05T
$0.56 \pm 0.41 \pm 0.16$	<sup>2</sup> ABE	03C BELL	Repl. by ABE 03H
$0.15 \pm 0.29 \pm 0.07$	<sup>2</sup> ABE	03H BELL	Repl. by CHEN 05B

<sup>1</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0 K^+ K^-$  decay.

<sup>2</sup> Quotes  $A_{\phi K_S^0}$  which is equal to  $-C_{\phi K_S^0}$ .

<sup>3</sup> Uses Dalitz plot analysis of  $B^0 \rightarrow K_S^0 K^+ K^-$  decays and the first of four consistent solutions that may be preferred.

<sup>4</sup> Result combines  $B$ -meson final states  $\phi K_S^0$  and  $\phi K_L^0$  by assuming  $S_{\phi K_S^0} = -S_{\phi K_L^0}$ .

 $S_{\phi K_S^0} (B^0 \rightarrow \phi K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.59 \pm 0.14</math> OUR AVERAGE</b>			

$0.66 \pm 0.17 \pm 0.07$	<sup>1</sup> LEES	12o BABR	$e^+e^- \rightarrow \gamma(4S)$
$0.50 \pm 0.21 \pm 0.06$	<sup>2</sup> CHEN	07 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.21 \pm 0.26 \pm 0.11$	<sup>2,3</sup> AUBERT	07AX BABR	Repl. by LEES 12o
$0.50 \pm 0.25 \pm 0.07 \pm 0.04$	<sup>2</sup> AUBERT	05T BABR	Repl. by AUBERT 07AX
$0.08 \pm 0.33 \pm 0.09$	<sup>2</sup> CHEN	05B BELL	Repl. by CHEN 07
$0.47 \pm 0.34 \pm 0.08 \pm 0.06$	<sup>2</sup> AUBERT,B	04G BABR	Repl. by AUBERT 05T
$-0.73 \pm 0.64 \pm 0.22$	ABE	03C BELL	Repl. by ABE 03H
$-0.96 \pm 0.50 \pm 0.09 \pm 0.11$	ABE	03H BELL	Repl. by CHEN 05B

<sup>1</sup> Uses Dalitz plot analysis of the  $B^0 \rightarrow K_S^0 K^+ K^-$  decay.

<sup>2</sup> Result combines  $B$ -meson final states  $\phi K_S^0$  and  $\phi K_L^0$  by assuming  $S_{\phi K_S^0} = -S_{\phi K_L^0}$ .

<sup>3</sup> Reports  $\beta_{eff}$ . We quote  $S$  obtained from epaps: E-PRLTAO-99-076741.

 $C_{K_S K_S K_S} (B^0 \rightarrow K_S K_S K_S)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.23 \pm 0.14</math> OUR AVERAGE</b>			

$-0.17 \pm 0.18 \pm 0.04$	LEES	12i BABR	$e^+e^- \rightarrow \gamma(4S)$
$-0.31 \pm 0.20 \pm 0.07$	<sup>1</sup> CHEN	07 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.02 \pm 0.21 \pm 0.05$	AUBERT	07AT BABR	Repl. by LEES 12i
$-0.34 \pm 0.28 \pm 0.05$	AUBERT,B	05 BABR	Repl. by AUBERT 07AT
$-0.54 \pm 0.34 \pm 0.09$	<sup>1</sup> SUMISAWA	05 BELL	Repl. by CHEN 07

<sup>1</sup> Belle Collab. quotes  $A_{K_S K_S K_S}$  which is equal to  $-C_{K_S K_S K_S}$ .

 $S_{K_S K_S K_S} (B^0 \rightarrow K_S K_S K_S)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.5 \pm 0.6</math> OUR AVERAGE</b>			Error includes scale factor of 3.0.

$-0.94 \pm 0.24 \pm 0.21$	LEES	12i BABR	$e^+e^- \rightarrow \gamma(4S)$
$0.30 \pm 0.32 \pm 0.08$	CHEN	07 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.71 \pm 0.24 \pm 0.04$	AUBERT	07AT BABR	Repl. by LEES 12i
$-0.71 \pm 0.38 \pm 0.32$	AUBERT,B	05 BABR	Repl. by AUBERT 07AT
$1.26 \pm 0.68 \pm 0.20$	SUMISAWA	05 BELL	Repl. by CHEN 07.

 $C_{K_S^0 \pi^0 \gamma} (B^0 \rightarrow K_S^0 \pi^0 \gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.36 \pm 0.33 \pm 0.04</math></b>	<sup>1</sup> AUBERT	08BA BABR	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.20 \pm 0.20 \pm 0.06$	<sup>2,3</sup> USHIRODA	06 BELL	$e^+e^- \rightarrow \gamma(4S)$
$-1.0 \pm 0.5 \pm 0.2$	<sup>1</sup> AUBERT,B	05P BABR	Repl. by AUBERT 08BA
$-0.03 \pm 0.34 \pm 0.11$	<sup>3</sup> USHIRODA	05 BELL	Repl. by USHIRODA 06

<sup>1</sup> Requires  $1.1 < M_{K_S^0 \pi^0} < 1.8 \text{ GeV}/c^2$ .

<sup>2</sup> Requires  $M_{K_S^0 \pi^0} < 1.8 \text{ GeV}/c^2$ .

<sup>3</sup> Reports  $A_{K_S^0 \pi^0 \gamma}$ , which is  $-C_{K_S^0 \pi^0 \gamma}$ .

 $S_{K_S^0 \pi^0 \gamma} (B^0 \rightarrow K_S^0 \pi^0 \gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.78 \pm 0.59 \pm 0.09</math></b>	<sup>1</sup> AUBERT	08BA BABR	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.10 \pm 0.31 \pm 0.07$	<sup>2</sup> USHIRODA	06 BELL	$e^+e^- \rightarrow \gamma(4S)$
$0.9 \pm 1.0 \pm 0.2$	<sup>1</sup> AUBERT,B	05P BABR	Repl. by AUBERT 08BA
$-0.58 \pm 0.46 \pm 0.11$	USHIRODA	05 BELL	Repl. by USHIRODA 06

<sup>1</sup> Requires  $1.1 < M_{K_S^0 \pi^0} < 1.8 \text{ GeV}/c^2$ .

<sup>2</sup> Requires  $M_{K_S^0 \pi^0} < 1.8 \text{ GeV}/c^2$ .

 $C_{K^*(892)^0 \gamma} (B^0 \rightarrow K^*(892)^0 \gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.04 \pm 0.16</math> OUR AVERAGE</b>			Error includes scale factor of 1.2.

$-0.14 \pm 0.16 \pm 0.03$	<sup>1</sup> AUBERT	08BA BABR	$e^+e^- \rightarrow \gamma(4S)$
$0.20 \pm 0.24 \pm 0.05$	<sup>1,2</sup> USHIRODA	06 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.40 \pm 0.23 \pm 0.03$	AUBERT,B	05P BABR	Repl. by AUBERT 08BA
$-0.57 \pm 0.32 \pm 0.09$	<sup>3</sup> AUBERT,B	04Z BABR	Repl. by AUBERT,B 05P

<sup>1</sup> Requires  $0.8 < M_{K_S^0 \pi^0} < 1.0 \text{ GeV}/c^2$ .

<sup>2</sup> Reports value of  $A$  which is equal to  $-C$ .

<sup>3</sup> Based on a total signal of  $105 \pm 14$  events with  $K^*(892)^0 \rightarrow K_S^0 \pi^0$  only.

 $S_{K^*(892)^0 \gamma} (B^0 \rightarrow K^*(892)^0 \gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.15 \pm 0.22</math> OUR AVERAGE</b>			

$-0.03 \pm 0.29 \pm 0.03$	<sup>1</sup> AUBERT	08BA BABR	$e^+e^- \rightarrow \gamma(4S)$
$-0.32 \pm 0.36 \pm 0.33$	<sup>1</sup> USHIRODA	06 BELL	$e^+e^- \rightarrow \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.21 \pm 0.40 \pm 0.05$	AUBERT,B	05P BABR	Repl. by AUBERT 08BA
$-0.79 \pm 0.63 \pm 0.50$	<sup>2</sup> USHIRODA	05 BELL	Repl. by USHIRODA 06
$0.25 \pm 0.63 \pm 0.14$	<sup>3</sup> AUBERT,B	04Z BABR	Repl. by AUBERT,B 05P

<sup>1</sup> Requires  $0.8 < M_{K_S^0 \pi^0} < 1.0 \text{ GeV}/c^2$ .

<sup>2</sup> Assumes  $C(B^0 \rightarrow K^*(892)^0 \gamma) = 0$ .

<sup>3</sup> Based on a total signal of  $105 \pm 14$  events with  $K^*(892)^0 \rightarrow K_S^0 \pi^0$  only.

 $C_{\eta K^0 \gamma} (B^0 \rightarrow \eta K^0 \gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.32 \pm 0.40 \pm 0.39</math> OUR AVERAGE</b>			

$-0.32 \pm 0.40 \pm 0.39$	<sup>1</sup> AUBERT	09 BABR	$e^+e^- \rightarrow \gamma(4S)$
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<sup>1</sup>  $m_{\eta K} < 3.25 \text{ GeV}/c^2$ .

 $S_{\eta K^0 \gamma} (B^0 \rightarrow \eta K^0 \gamma)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.18 \pm 0.49 \pm 0.46</math> OUR AVERAGE</b>			

$-0.18 \pm 0.49 \pm 0.46$	<sup>1</sup> AUBERT	09 BABR	$e^+e^- \rightarrow \gamma(4S)$
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<sup>1</sup>  $m_{\eta K} < 3.25 \text{ GeV}/c^2$ .

# Meson Particle Listings

$B^0$

## $C_{K^0\phi\gamma}(B^0 \rightarrow K^0\phi\gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.35 \pm 0.56^{+0.10}_{-0.23}$	<sup>1</sup> SAHOO	11A	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Reports value of $A$ , which is equal to $-C$ .			

## $S_{K^0\phi\gamma}(B^0 \rightarrow K^0\phi\gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.74^{+0.72+0.10}_{-1.05-0.24}$	SAHOO	11A	BELL $e^+e^- \rightarrow \Upsilon(4S)$

## $C(B^0 \rightarrow K_S^0\rho^0\gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.05 \pm 0.18 \pm 0.06$	<sup>1,2</sup> LI	08F	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Requires $M_{K_S^0\pi^+\pi^-} < 1.8 \text{ GeV}/c^2$ and $0.6 < M_{\pi^+\pi^-} < 0.9 \text{ GeV}/c^2$ .			
<sup>2</sup> Reports value of $A_{\text{eff}}$ which is equal to $-C$ , and includes the non-resonant $\pi^+\pi^-$ contribution in the $\rho^0$ region.			

## $S(B^0 \rightarrow K_S^0\rho^0\gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.11 \pm 0.33^{+0.05}_{-0.09}$	<sup>1</sup> LI	08F	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Requires $M_{K_S^0\pi^+\pi^-} < 1.8 \text{ GeV}/c^2$ .			

## $C(B^0 \rightarrow \rho^0\gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.44 \pm 0.49 \pm 0.14$	<sup>1</sup> USHIRODA	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Reports value of $A$ which is equal to $-C$ .			

## $S(B^0 \rightarrow \rho^0\gamma)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.83 \pm 0.65 \pm 0.18$	USHIRODA	08	BELL $e^+e^- \rightarrow \Upsilon(4S)$

## $C_{\pi\pi}(B^0 \rightarrow \pi^+\pi^-)$

$C_{\pi\pi}$  is defined as  $(1-|\lambda|^2)/(1+|\lambda|^2)$ , where the quantity  $\lambda=q/p \bar{A}_f/A_f$  is a phase convention independent observable quantity for the final state  $f$ . For details, see the review on “CP Violation” in the Reviews section.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.31 \pm 0.05</math> OUR AVERAGE</b>			
$-0.38 \pm 0.15 \pm 0.02$	AAIJ	13B0	LHCB $pp$ at 7 TeV
$-0.33 \pm 0.06 \pm 0.03$	<sup>1</sup> DALSENO	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.25 \pm 0.08 \pm 0.02$	LEES	13D	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.21 \pm 0.09 \pm 0.02$	AUBERT	07AF	BABR Repl. by LEES 13D
$-0.55 \pm 0.08 \pm 0.05$	<sup>1</sup> ISHINO	07	BELL Repl. by DALSENO 13
$-0.56 \pm 0.12 \pm 0.06$	<sup>1</sup> ABE	05D	BELL Repl. by ISHINO 07
$-0.09 \pm 0.15 \pm 0.04$	AUBERT,BE	05	BABR Repl. by AUBERT 07AF
$-0.58 \pm 0.15 \pm 0.07$	<sup>1</sup> ABE	04E	BELL Repl. by ABE 05D
$-0.77 \pm 0.27 \pm 0.08$	<sup>1</sup> ABE	03G	BELL Repl. by ABE 04E.
$-0.94^{+0.31}_{-0.25} \pm 0.09$	<sup>1</sup> ABE	02M	BELL Repl. by ABE 03G
$-0.25^{+0.45}_{-0.47} \pm 0.14$	<sup>2</sup> AUBERT	02D	BABR Repl. by AUBERT 02Q
$-0.30 \pm 0.25 \pm 0.04$	<sup>3</sup> AUBERT	02Q	BABR Repl. by AUBERT,BE 05
<sup>1</sup> Paper reports $A_{\pi\pi}$ which equals to $-C_{\pi\pi}$ .			
<sup>2</sup> Corresponds to 90% confidence range $-1.0 < C_{\pi\pi} < 0.47$ .			
<sup>3</sup> Corresponds to 90% confidence range $-0.72 < C_{\pi\pi} < 0.12$ .			

## $S_{\pi\pi}(B^0 \rightarrow \pi^+\pi^-)$

$S_{\pi\pi} = 2\text{Im}\lambda/(1+|\lambda|^2)$ , see the note in the  $C_{\pi\pi}$  datablock above.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.67 \pm 0.06</math> OUR AVERAGE</b>			
$-0.71 \pm 0.13 \pm 0.02$	AAIJ	13B0	LHCB $pp$ at 7 TeV
$-0.64 \pm 0.08 \pm 0.03$	<sup>1</sup> DALSENO	13	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.68 \pm 0.10 \pm 0.03$	LEES	13D	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.60 \pm 0.11 \pm 0.03$	AUBERT	07AF	BABR Repl. by LEES 13D
$-0.61 \pm 0.10 \pm 0.04$	ISHINO	07	BELL Repl. by DALSENO 13
$-0.67 \pm 0.16 \pm 0.06$	<sup>2</sup> ABE	05D	BELL Repl. by ISHINO 07
$-0.30 \pm 0.17 \pm 0.03$	AUBERT,BE	05	BABR Repl. by AUBERT 07AF
$-1.00 \pm 0.21 \pm 0.07$	<sup>3</sup> ABE	04E	BELL Repl. by ABE 05D
$-1.23 \pm 0.41^{+0.08}_{-0.07}$	ABE	03G	BELL Repl. by ABE 04E.
$-1.21^{+0.38+0.16}_{-0.27-0.13}$	ABE	02M	BELL Repl. by ABE 03G
$0.03^{+0.52}_{-0.56} \pm 0.11$	<sup>4</sup> AUBERT	02D	BABR Repl. by AUBERT 02Q
$0.02 \pm 0.34 \pm 0.05$	<sup>5</sup> AUBERT	02Q	BABR Repl. by AUBERT,BE 05

<sup>1</sup> An isospin analysis using other BELLE measurements, disfavors the region of  $23.8^\circ < \phi_2 < 66.8^\circ$  at 68% CL.  
<sup>2</sup> Rule out the  $CP$ -conserving case,  $C_{\pi\pi} = S_{\pi\pi} = 0$ , at the 5.4 sigma level.  
<sup>3</sup> Rule out the  $CP$ -conserving case,  $C_{\pi\pi} = S_{\pi\pi} = 0$ , at the 5.2 sigma level.  
<sup>4</sup> Corresponds to 90% confidence range  $-0.89 < S_{\pi\pi} < 0.85$ .  
<sup>5</sup> Corresponds to 90% confidence range  $-0.54 < S_{\pi\pi} < 0.58$ .

## $C_{\pi^0\pi^0}(B^0 \rightarrow \pi^0\pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.43 \pm 0.24</math> OUR AVERAGE</b>			
$-0.43 \pm 0.26 \pm 0.05$	LEES	13D	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.44^{+0.52}_{-0.53} \pm 0.17$	<sup>1</sup> CHAO	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.49 \pm 0.35 \pm 0.05$	AUBERT	07Bc	BABR Repl. by LEES 13D
$-0.12 \pm 0.56 \pm 0.06$	<sup>2</sup> AUBERT	05L	BABR Repl. by AUBERT 07Bc
<sup>1</sup> BELLE Collab. quotes $A_{\pi^0\pi^0}$ which is equal to $-C_{\pi^0\pi^0}$ .			
<sup>2</sup> Corresponds to a 90% CL interval of $-0.88 < A_{CP} < 0.64$ .			

## $C_{\rho\pi}(B^0 \rightarrow \rho^+\pi^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.03 \pm 0.07</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.		
$0.016 \pm 0.059 \pm 0.036$	<sup>1</sup> LEES	13J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.13 \pm 0.09 \pm 0.05$	<sup>1</sup> KUSAKA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.15 \pm 0.09 \pm 0.05$	AUBERT	07AA	BABR Repl. by LEES 13J
$0.25 \pm 0.17^{+0.02}_{-0.06}$	WANG	05	BELL Repl. by KUSAKA 07
$0.36 \pm 0.18 \pm 0.04$	AUBERT	03T	BABR Repl. by AUBERT 07AA
<sup>1</sup> Uses time-dependent Dalitz plot analysis of $B^0 \rightarrow \pi^+\pi^-\pi^0$ decays.			

## $S_{\rho\pi}(B^0 \rightarrow \rho^+\pi^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.05 \pm 0.07</math> OUR AVERAGE</b>			
$0.053 \pm 0.081 \pm 0.034$	<sup>1</sup> LEES	13J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.06 \pm 0.13 \pm 0.05$	<sup>1</sup> KUSAKA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.03 \pm 0.11 \pm 0.04$	AUBERT	07AA	BABR Repl. by LEES 13J
$-0.28 \pm 0.23^{+0.10}_{-0.08}$	WANG	05	BELL Repl. by KUSAKA 07
$0.19 \pm 0.24 \pm 0.03$	AUBERT	03T	BABR Repl. by AUBERT 07AA
<sup>1</sup> Uses time-dependent Dalitz plot analysis of $B^0 \rightarrow \pi^+\pi^-\pi^0$ decays.			

## $\Delta C_{\rho\pi}(B^0 \rightarrow \rho^+\pi^-)$

$\Delta C_{\rho\pi}$  describes the asymmetry between the rates  $\Gamma(B^0 \rightarrow \rho^+\pi^-) + \Gamma(\bar{B}^0 \rightarrow \rho^-\pi^+)$  and  $\Gamma(B^0 \rightarrow \rho^-\pi^+) + \Gamma(\bar{B}^0 \rightarrow \rho^+\pi^-)$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.27 \pm 0.06</math> OUR AVERAGE</b>			
$0.234 \pm 0.061 \pm 0.048$	<sup>1</sup> LEES	13J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.36 \pm 0.10 \pm 0.05$	<sup>1</sup> KUSAKA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.39 \pm 0.09 \pm 0.09$	AUBERT	07AA	BABR Repl. by LEES 13J
$0.38 \pm 0.18^{+0.02}_{-0.04}$	WANG	05	BELL Repl. by KUSAKA 07
$0.28^{+0.18}_{-0.19} \pm 0.04$	AUBERT	03T	BABR Repl. by AUBERT 07AA
<sup>1</sup> Uses time-dependent Dalitz plot analysis of $B^0 \rightarrow \pi^+\pi^-\pi^0$ decays.			

## $\Delta S_{\rho\pi}(B^0 \rightarrow \rho^+\pi^-)$

$\Delta S_{\rho\pi}$  is related to the strong phase difference between the amplitudes contributing to  $B^0 \rightarrow \rho^+\pi^-$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.01 \pm 0.08</math> OUR AVERAGE</b>			
$0.054 \pm 0.082 \pm 0.039$	<sup>1</sup> LEES	13J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.08 \pm 0.13 \pm 0.05$	<sup>1</sup> KUSAKA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.01 \pm 0.14 \pm 0.06$	AUBERT	07AA	BABR Repl. by LEES 13J
$-0.30 \pm 0.24 \pm 0.09$	WANG	05	BELL Repl. by KUSAKA 07
$0.15 \pm 0.25 \pm 0.03$	AUBERT	03T	BABR Repl. by AUBERT 07AA
<sup>1</sup> Uses time-dependent Dalitz plot analysis of $B^0 \rightarrow \pi^+\pi^-\pi^0$ decays.			

## $C_{\rho^0\pi^0}(B^0 \rightarrow \rho^0\pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.27 \pm 0.24</math> OUR AVERAGE</b>			
$0.19 \pm 0.23 \pm 0.15$	<sup>1</sup> LEES	13J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.49 \pm 0.36 \pm 0.28$	<sup>1,2</sup> KUSAKA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.10 \pm 0.40 \pm 0.53$	AUBERT	07AA	BABR Repl. by LEES 13J
$0.53 \pm 0.67^{+0.10}_{-0.84-0.15}$	<sup>2</sup> DRAGIC	06	BELL Repl. by KUSAKA 07
<sup>1</sup> Uses time-dependent Dalitz plot analysis of $B^0 \rightarrow \pi^+\pi^-\pi^0$ decays.			
<sup>2</sup> Quotes $A_{\rho^0\pi^0}$ which is equal to $-C_{\rho^0\pi^0}$ .			

## $S_{\rho^0\pi^0}(B^0 \rightarrow \rho^0\pi^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.23 \pm 0.34</math> OUR AVERAGE</b>			
$-0.37 \pm 0.34 \pm 0.20$	<sup>1</sup> LEES	13J	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.17 \pm 0.57 \pm 0.35$	<sup>1</sup> KUSAKA	07	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.04 \pm 0.44 \pm 0.18$	AUBERT	07AA	BABR Repl. by LEES 13J
<sup>1</sup> Uses time-dependent Dalitz plot analysis of $B^0 \rightarrow \pi^+\pi^-\pi^0$ decays.			

$C_{a_1\pi}(B^0 \rightarrow a_1(1260)^+\pi^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.05 \pm 0.11</math> OUR AVERAGE</b>			
$-0.01 \pm 0.11 \pm 0.09$	DALSENO	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.10 \pm 0.15 \pm 0.09$	AUBERT	07o	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $S_{a_1\pi}(B^0 \rightarrow a_1(1260)^+\pi^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.2 \pm 0.4</math> OUR AVERAGE</b>			Error includes scale factor of 3.2.
$-0.51 \pm 0.14 \pm 0.08$	DALSENO	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.37 \pm 0.21 \pm 0.07$	AUBERT	07o	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $\Delta C_{a_1\pi}(B^0 \rightarrow a_1(1260)^+\pi^-)$ 

$\Delta C_{a_1\pi}$  describes the asymmetry between the rates  $\Gamma(B^0 \rightarrow a_1^+\pi^-) + \Gamma(\bar{B}^0 \rightarrow a_1^-\pi^+)$  and  $\Gamma(B^0 \rightarrow a_1^-\pi^+) + \Gamma(\bar{B}^0 \rightarrow a_1^+\pi^-)$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.43 \pm 0.14</math> OUR AVERAGE</b>			Error includes scale factor of 1.3.
$0.54 \pm 0.11 \pm 0.07$	DALSENO	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.26 \pm 0.15 \pm 0.07$	AUBERT	07o	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $\Delta S_{a_1\pi}(B^0 \rightarrow a_1(1260)^+\pi^-)$ 

$\Delta S_{a_1\pi}$  is related to the strong phase difference between the amplitudes contributing to  $B^0 \rightarrow a_1\pi$  decays.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.11 \pm 0.12</math> OUR AVERAGE</b>			
$-0.09 \pm 0.14 \pm 0.06$	DALSENO	12	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.14 \pm 0.21 \pm 0.06$	AUBERT	07o	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $C(B^0 \rightarrow b_1^-K^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.22 \pm 0.23 \pm 0.05</math></b>			
	AUBERT	07Bi	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $\Delta C(B^0 \rightarrow b_1^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-1.04 \pm 0.23 \pm 0.08</math></b>			
	AUBERT	07Bi	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $C_{\rho^0\rho}(B^0 \rightarrow \rho^0\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.2 \pm 0.8 \pm 0.3</math></b>			
	AUBERT	08Bb	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $S_{\rho^0\rho}(B^0 \rightarrow \rho^0\rho^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.3 \pm 0.7 \pm 0.2</math></b>			
	AUBERT	08Bb	BABR $e^+e^- \rightarrow \Upsilon(4S)$

 $C_{\rho\rho}(B^0 \rightarrow \rho^+\rho^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.09</math> OUR AVERAGE</b>			
$0.00 \pm 0.10 \pm 0.06$	1 VANHOEFER	16	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.01 \pm 0.15 \pm 0.06$	AUBERT	07Bf	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.16 \pm 0.21 \pm 0.08$	1 SOMOV	07	BELL Repl. by VANHOEFER 16
$-0.00 \pm 0.30 \pm 0.09$	1 SOMOV	06	BELL Repl. by SOMOV 07
$-0.03 \pm 0.18 \pm 0.09$	AUBERT,B	05c	BABR Repl. by AUBERT 07Bf
$-0.17 \pm 0.27 \pm 0.14$	AUBERT,B	04R	BABR Repl. by AUBERT,B 05c
1 BELLE Collab. quotes $A_{CP}$ which is equal to $-C$ .			

 $S_{\rho\rho}(B^0 \rightarrow \rho^+\rho^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.14 \pm 0.13</math> OUR AVERAGE</b>			
$-0.13 \pm 0.15 \pm 0.05$	VANHOEFER	16	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.17 \pm 0.20 \pm 0.05$	AUBERT	07Bf	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.19 \pm 0.30 \pm 0.08$	SOMOV	07	BELL Repl. by VANHOEFER 16
$0.08 \pm 0.41 \pm 0.09$	SOMOV	06	BELL Repl. by SOMOV 07
$-0.33 \pm 0.24 \pm 0.08$	AUBERT,B	05c	BABR Repl. by AUBERT 07Bf
$-0.42 \pm 0.42 \pm 0.14$	AUBERT,B	04R	BABR Repl. by AUBERT,B 05c

 $|\lambda|(B^0 \rightarrow J/\psi K^*(892)^0)$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.25</math></b>	95	1 AUBERT,B	04H	BABR $e^+e^- \rightarrow \Upsilon(4S)$

1 Uses the measured cosine coefficients  $C$  and  $\bar{C}$  and assumes  $|q/p| = 1$ .

 $\cos 2\beta(B^0 \rightarrow J/\psi K^*(892)^0)$ 

$\beta(\phi_1)$  is one of the angles of CKM unitarity triangle, see the review on "CP" Violation in the Reviews section.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.7 \pm 0.7</math> OUR AVERAGE</b>			Error includes scale factor of 1.6.
$2.72 \pm 0.50 \pm 0.79$	1 AUBERT	05P	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.87 \pm 0.74 \pm 0.12$	2 ITOH	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

1 The measurement is obtained when  $\sin 2\beta$  is fixed to 0.726 and the sign of  $\cos 2\beta$  is positive with 86% confidence level.

2 The measurement is obtained with  $\sin 2\beta$  fixed to 0.731.

 $\cos 2\beta(B^0 \rightarrow [K_S^0\pi^+\pi^-]_{D^{(*)}}h^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.0 \pm 0.6</math> OUR AVERAGE</b>			Error includes scale factor of 1.8.
$0.42 \pm 0.49 \pm 0.16$	1 AUBERT	07Bh	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$1.87 \pm 0.40 \pm 0.22$	2 KROKOVNY	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.53 \pm 0.32$			

1 AUBERT 07Bh evaluates the likelihoods for the positive and negative solutions assuming  $\sin(2\beta_{eff}) = 0.678$ . It quotes  $L_+/L_- = 0.86$  corresponding to a likelihood ratio of  $L_+/L_- = 6.14$  in favor of the positive solution.

2 KROKOVNY 06 evaluates the likelihoods for the positive and negative solutions assuming  $\sin(2\beta_{eff}) = 0.689$ . It quotes  $L_+/(L_+ + L_-) = 0.983$  corresponding to a likelihood ratio of  $L_+/L_- = 57.8$  in favor of the positive solution.

 $(S_+ + S_-)/2(B^0 \rightarrow D^{*-}\pi^+)$ 

$S_{\pm} = -\frac{2Im(\lambda_{\pm})}{1+|\lambda_{\pm}|^2}$  where  $\lambda_+$  and  $\lambda_-$  are defined in the  $C_{\pi\pi}$  datablock above for  $B^0 \rightarrow D^{*-}\pi^+$  and  $\bar{B}^0 \rightarrow D^{*+}\pi^-$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.039 \pm 0.011</math> OUR AVERAGE</b>			
$-0.046 \pm 0.013 \pm 0.015$	1 BAHINIPATI	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$-0.040 \pm 0.023 \pm 0.010$	2 AUBERT	06Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.034 \pm 0.014 \pm 0.009$	1 AUBERT	05Z	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.039 \pm 0.020 \pm 0.013$	3 RONGA	06	BELL Repl. by BAHINIPATI 11
$-0.030 \pm 0.028 \pm 0.018$	1 GERSHON	05	BELL Repl. by RONGA 06
$-0.068 \pm 0.038 \pm 0.020$	2 AUBERT	04V	BABR Repl. by AUBERT 06Y
$-0.063 \pm 0.024 \pm 0.014$	1 AUBERT	04W	BABR Repl. by AUBERT 05Z
$0.060 \pm 0.040 \pm 0.019$	2 SARANGI	04	BELL Repl. by RONGA 06

1 Uses partially reconstructed  $B^0 \rightarrow D^{*\pm}\pi^\mp$  decays.

2 Uses fully reconstructed  $B^0 \rightarrow D^{*\pm}\pi^\mp$  decays.

3 Combines the results from fully reconstructed and partially reconstructed  $D^*\pi$  events by taking weighted averages. Assumes that systematic errors from physics parameters and fit biases in the two measurements are 100% correlated.

 $(S_- - S_+)/2(B^0 \rightarrow D^{*-}\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.009 \pm 0.015</math> OUR AVERAGE</b>			
$-0.015 \pm 0.013 \pm 0.015$	1 BAHINIPATI	11	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.049 \pm 0.042 \pm 0.015$	2 AUBERT	06Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.019 \pm 0.022 \pm 0.013$	1 AUBERT	05Z	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.011 \pm 0.020 \pm 0.013$	3 RONGA	06	BELL Repl. by BAHINIPATI 11
$-0.005 \pm 0.028 \pm 0.018$	1 GERSHON	05	BELL Repl. by RONGA 06
$0.031 \pm 0.070 \pm 0.033$	2 AUBERT	04V	BABR Repl. by AUBERT 06Y
$-0.004 \pm 0.037 \pm 0.014$	1 AUBERT	04W	BABR Repl. by AUBERT 05Z
$0.049 \pm 0.040 \pm 0.019$	2 SARANGI	04	BELL Repl. by RONGA 06

1 Uses partially reconstructed  $B^0 \rightarrow D^{*\pm}\pi^\mp$  decays.

2 Uses fully reconstructed  $B^0 \rightarrow D^{*\pm}\pi^\mp$  decays.

3 Combines the results from fully reconstructed and partially reconstructed  $D^*\pi$  events by taking weighted averages. Assumes that systematic errors from physics parameters and fit biases in the two measurements are 100% correlated.

 $(S_+ + S_-)/2(B^0 \rightarrow D^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.046 \pm 0.023</math> OUR AVERAGE</b>			
$-0.010 \pm 0.023 \pm 0.07$	1 AUBERT	06Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.050 \pm 0.021 \pm 0.012$	2 RONGA	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.022 \pm 0.038 \pm 0.020$	1 AUBERT	04V	BABR Repl. by AUBERT 06Y
$-0.062 \pm 0.037 \pm 0.018$	1 SARANGI	04	BELL Repl. by RONGA 06

1 Uses fully reconstructed  $B^0 \rightarrow D^\pm\pi^\mp$  decays.

2 Combines the results from fully reconstructed and partially reconstructed  $D\pi$  events by taking weighted averages. Assumes that systematic errors from physics parameters and fit biases in the two measurements are 100% correlated.

 $(S_- - S_+)/2(B^0 \rightarrow D^-\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.022 \pm 0.021</math> OUR AVERAGE</b>			
$-0.033 \pm 0.042 \pm 0.012$	1 AUBERT	06Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$-0.019 \pm 0.021 \pm 0.012$	2 RONGA	06	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.025 \pm 0.068 \pm 0.033$	1 AUBERT	04V	BABR Repl. by AUBERT 06Y
$-0.025 \pm 0.037 \pm 0.018$	1 SARANGI	04	BELL Repl. by RONGA 06

1 Uses fully reconstructed  $B^0 \rightarrow D^\pm\pi^\mp$  decays.

2 Combines the results from fully reconstructed and partially reconstructed  $D\pi$  events by taking weighted averages. Assumes that systematic errors from physics parameters and fit biases in the two measurements are 100% correlated.

 $(S_+ + S_-)/2(B^0 \rightarrow D^-\rho^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.024 \pm 0.031 \pm 0.009</math></b>			
	1 AUBERT	06Y	BABR $e^+e^- \rightarrow \Upsilon(4S)$

1 Uses fully reconstructed  $B^0 \rightarrow D^-\rho^+$  decays.

# Meson Particle Listings

## $B^0$

### $(S_- - S_+)/2 (B^0 \rightarrow D^- \rho^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.098 \pm 0.055 \pm 0.018$	<sup>1</sup> AUBERT	06Y	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses fully reconstructed $B^0 \rightarrow D^- \rho^+$ decays.			

### $C_{\eta_c K_S^0} (B^0 \rightarrow \eta_c K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.080 \pm 0.124 \pm 0.029$	AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

### $S_{\eta_c K_S^0} (B^0 \rightarrow \eta_c K_S^0)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.925 \pm 0.160 \pm 0.057$	AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$

### $C_{c\bar{c}K^{(*)}} (B^0 \rightarrow c\bar{c}K^{(*)0})$

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.5 \pm 1.7</math> OUR EVALUATION</b>			
<b><math>0.5 \pm 1.6</math> OUR AVERAGE</b>			
$-0.6 \pm 1.6 \pm 1.2$	<sup>1</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-29 \begin{smallmatrix} +53 \\ -44 \end{smallmatrix} \pm 6$	<sup>2</sup> AUBERT	09AU	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$-2.4 \pm 2.0 \pm 1.6$	<sup>3</sup> AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-4 \pm 7 \pm 5$	<sup>4</sup> SAHOO	08	BELL Repl. by ADACHI 12A
$4.9 \pm 2.3 \pm 1.8$	<sup>3</sup> AUBERT	07AY	BABR Repl. by AUBERT 09K
$-1.8 \pm 2.1 \pm 1.4$	<sup>5</sup> CHEN	07	BELL Repl. by ADACHI 12A
$-0.7 \pm 4.1 \pm 3.3$	<sup>6</sup> ABE	05B	BELL Repl. by CHEN 07
$5.1 \pm 3.2 \pm 1.4$	<sup>7</sup> AUBERT	05F	BABR Repl. by AUBERT 07AY
$5.1 \pm 5.1 \pm 2.6$	<sup>8</sup> ABE	02Z	BELL Repl. by ABE 05B
$5.3 \pm 5.4 \pm 3.2$	<sup>9</sup> AUBERT	02P	BABR Repl. by AUBERT 05F

- Measurement based on  $B^0 \rightarrow J/\psi K_S^0$ ,  $B^0 \rightarrow \psi(2S) K_S^0$ ,  $B^0 \rightarrow J/\psi K_L^0$ , and  $B^0 \rightarrow \chi_{c1}(1P) K_S^0$  decays.
- Uses Dalitz plot analysis of  $B^0 \rightarrow K^0 \pi^+ \pi^-$  decays and the first of two equivalent solutions is used.
- Measurement based on  $B^0 \rightarrow c\bar{c}K^{(*)0}$  decays.
- Reports value of A of  $B^0 \rightarrow \psi(2S) K^0$  which is equal to  $-C$ .
- Reports value of A of  $B^0 \rightarrow J/\psi K^0$  which is equal to  $-C$ .
- Measurement based on  $152 \times 10^6 B\bar{B}$  pairs.
- Measurement based on  $227 \times 10^6 B\bar{B}$  pairs.
- Measured with both  $\eta_f = \pm 1$  samples.
- Measured with the high purity of  $\eta_f = -1$  samples.

### $\sin(2\beta)$

For a discussion of  $CP$  violation, see the review on “ $CP$  Violation” in the Reviews section.  $\sin(2\beta)$  is a measure of the  $CP$ -violating amplitude in the  $B_d^0 \rightarrow J/\psi(1S) K_S^0$ .

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.679 \pm 0.020</math> OUR EVALUATION</b>			
<b><math>0.677 \pm 0.020</math> OUR AVERAGE</b>			
$0.667 \pm 0.023 \pm 0.012$	<sup>1</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.57 \pm 0.58 \pm 0.06$	<sup>2</sup> SATO	12	BELL $e^+ e^- \rightarrow \Upsilon(5S)$
$0.69 \pm 0.52 \pm 0.08$	<sup>3</sup> AUBERT	09AU	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.687 \pm 0.028 \pm 0.012$	<sup>4</sup> AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$1.56 \pm 0.42 \pm 0.21$	<sup>5</sup> AUBERT	04R	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.79 \begin{smallmatrix} +0.41 \\ -0.44 \end{smallmatrix}$	<sup>6</sup> AFFOLDER	00c	CDF $p\bar{p}$ at 1.8 TeV
$0.84 \begin{smallmatrix} +0.82 \\ -1.04 \end{smallmatrix} \pm 0.16$	<sup>7</sup> BARATE	00Q	ALEP $e^+ e^- \rightarrow Z$
$3.2 \begin{smallmatrix} +1.8 \\ -2.0 \end{smallmatrix} \pm 0.5$	<sup>8</sup> ACKERSTAFF	98z	OPAL $e^+ e^- \rightarrow Z$

• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.72 \pm 0.09 \pm 0.03$	<sup>9</sup> SAHOO	08	BELL Repl. by ADACHI 12A
$0.714 \pm 0.032 \pm 0.018$	<sup>4</sup> AUBERT	07AY	BABR Repl. by AUBERT 09K
$0.642 \pm 0.031 \pm 0.017$	CHEN	07	BELL Repl. by ADACHI 12A
$0.728 \pm 0.056 \pm 0.023$	<sup>10</sup> ABE	05B	BELL Repl. by CHEN 07
$0.722 \pm 0.040 \pm 0.023$	<sup>11</sup> AUBERT	05F	BABR Repl. by AUBERT 07AY
$0.99 \pm 0.14 \pm 0.06$	<sup>12</sup> ABE	02U	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.719 \pm 0.074 \pm 0.035$	<sup>13</sup> ABE	02Z	BELL Repl. by ABE 05B
$0.59 \pm 0.14 \pm 0.05$	<sup>14</sup> AUBERT	02N	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.741 \pm 0.067 \pm 0.034$	<sup>15</sup> AUBERT	02P	BABR Repl. by AUBERT 05F
$0.58 \begin{smallmatrix} +0.32 \\ -0.34 \end{smallmatrix} \pm 0.09$	ABASHIAN	01	BELL Repl. by ABE 01G
$0.99 \pm 0.14 \pm 0.06$	<sup>16</sup> ABE	01G	BELL Repl. by ABE 02Z
$0.34 \pm 0.20 \pm 0.05$	AUBERT	01	BABR Repl. by AUBERT 01B
$0.59 \pm 0.14 \pm 0.05$	<sup>16</sup> AUBERT	01B	BABR Repl. by AUBERT 02P
$1.8 \pm 1.1 \pm 0.3$	<sup>17</sup> ABE	98U	CDF Repl. by AFFOLDER 00c

- Measurement based on  $B^0 \rightarrow J/\psi K_S^0$ ,  $B^0 \rightarrow \psi(2S) K_S^0$ ,  $B^0 \rightarrow J/\psi K_L^0$ , and  $B^0 \rightarrow \chi_{c1}(1P) K_S^0$  decays.

- SATO 12 uses  $121 \text{ fb}^{-1}$  data collected on  $\Upsilon(5S)$  resonance. Uses the “ $B - \pi$  tagging” where  $B\pi^+$  and  $B\pi^-$  tagged  $J/\psi K_S^0$  events are compared.
- Uses Dalitz plot analysis of  $B^0 \rightarrow K^0 \pi^+ \pi^-$  decays and the first of two equivalent solutions.
- Measurement based on  $B^0 \rightarrow c\bar{c}K^{(*)0}$  decays.
- Measurement in which the  $J/\psi$  decays to hadrons or to muons that do not satisfy the standard identification criteria.
- AFFOLDER 00c uses about  $400 B^0 \rightarrow J/\psi(1S) K_S^0$  events. The production flavor of  $B^0$  was determined using three tagging algorithms: a same-side tag, a jet-charge tag, and a soft-lepton tag.
- BARATE 00Q uses 23 candidates for  $B^0 \rightarrow J/\psi(1S) K_S^0$  decays. A combination of jet-charge, vertex-charge, and same-side tagging techniques were used to determine the  $B^0$  production flavor.
- ACKERSTAFF 98z uses 24 candidates for  $B_d^0 \rightarrow J/\psi(1S) K_S^0$  decay. A combination of jet-charge and vertex-charge techniques were used to tag the  $B_d^0$  production flavor.
- Based on  $B^0 \rightarrow \psi(2S) K_S^0$  decays.
- Measurement based on  $152 \times 10^6 B\bar{B}$  pairs.
- Measurement based on  $227 \times 10^6 B\bar{B}$  pairs.
- ABE 02U result is based on the same analysis and data sample reported in ABE 01G.
- ABE 02Z result is based on  $85 \times 10^6 B\bar{B}$  pairs.
- AUBERT 02N result based on the same analysis and data sample reported in AUBERT 01B.
- AUBERT 02P result is based on  $88 \times 10^6 B\bar{B}$  pairs.
- First observation of  $CP$  violation in  $B^0$  meson system.
- ABE 98U uses  $198 \pm 17 B_d^0 \rightarrow J/\psi(1S) K^0$  events. The production flavor of  $B^0$  was determined using the same side tagging technique.

### $C_{J/\psi(nS) K^0} (B^0 \rightarrow J/\psi(nS) K^0)$

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.5 \pm 2.0</math> OUR EVALUATION</b>			
<b><math>-0.9 \pm 1.7</math> OUR AVERAGE</b>			Error includes scale factor of 1.1.
$-3.8 \pm 3.2 \pm 0.5$	<sup>1</sup> AAIJ	15N	LHCB $pp$ at 7, 8 TeV
$1.5 \pm 2.1 \begin{smallmatrix} +2.3 \\ -4.5 \end{smallmatrix}$	<sup>2,3</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-10.4 \pm 5.5 \begin{smallmatrix} +2.7 \\ -4.7 \end{smallmatrix}$	<sup>3,4</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$-1.9 \pm 2.6 \begin{smallmatrix} +4.1 \\ -1.7 \end{smallmatrix}$	<sup>3,5</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$8.9 \pm 7.6 \pm 2.0$	<sup>4</sup> AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$1.6 \pm 2.3 \pm 1.8$	AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3 \pm 9 \pm 1$	<sup>6</sup> AAIJ	13K	LHCB Repl. by AAIJ 15N
$-4 \pm 7 \pm 5$	<sup>3,4</sup> SAHOO	08	BELL Repl. by ADACHI 12A
$-1.8 \pm 2.1 \pm 1.4$	<sup>3</sup> CHEN	07	BELL Repl. by ADACHI 12A

- AAIJ 15N uses  $41,560$  flavor-tagged  $B_d \rightarrow J/\psi K_S^0$  events from  $3 \text{ fb}^{-1}$  of integrated luminosity. Provides the correlation coefficient  $\rho = 0.483$  between the statistical uncertainties of and measurements.
- Uses  $B^0 \rightarrow J/\psi K_S^0$  decays.
- The paper reports A, which is equal to  $-C$ .
- Uses  $B^0 \rightarrow \psi(2S) K_S^0$  decays.
- Uses  $B^0 \rightarrow J/\psi K_L^0$  decays.
- AAIJ 13K uses  $8200$  flavor-tagged  $B_d \rightarrow J/\psi K_S^0$  events from  $1 \text{ fb}^{-1}$  of integrated luminosity. Provides the correlation coefficient  $\rho = 0.42$  between the statistical uncertainties of  $S_{J/\psi(nS) K^0} (B^0 \rightarrow J/\psi(nS) K^0)$  and  $C_{J/\psi(nS) K^0} (B^0 \rightarrow J/\psi(nS) K^0)$  measurements.

### $S_{J/\psi(nS) K^0} (B^0 \rightarrow J/\psi(nS) K^0)$

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.676 \pm 0.021</math> OUR EVALUATION</b>			
<b><math>0.687 \pm 0.021</math> OUR AVERAGE</b>			Error includes scale factor of 1.2.
$0.731 \pm 0.035 \pm 0.020$	<sup>1</sup> AAIJ	15N	LHCB $pp$ at 7, 8 TeV
$0.670 \pm 0.029 \pm 0.013$	<sup>2</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.738 \pm 0.079 \pm 0.036$	<sup>3</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.642 \pm 0.047 \pm 0.021$	<sup>4</sup> ADACHI	12A	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
$0.57 \pm 0.58 \pm 0.06$	<sup>5</sup> SATO	12	BELL $e^+ e^- \rightarrow \Upsilon(5S)$
$0.897 \pm 0.100 \pm 0.036$	<sup>3</sup> AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.666 \pm 0.031 \pm 0.013$	AUBERT	09K	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
$0.79 \begin{smallmatrix} +0.41 \\ -0.44 \end{smallmatrix}$	<sup>6</sup> AFFOLDER	00c	CDF $p\bar{p}$ at 1.8 TeV
$0.84 \begin{smallmatrix} +0.82 \\ -1.04 \end{smallmatrix} \pm 0.16$	<sup>7</sup> BARATE	00Q	ALEP $e^+ e^- \rightarrow Z$
$3.2 \begin{smallmatrix} +1.8 \\ -2.0 \end{smallmatrix} \pm 0.5$	<sup>8</sup> ACKERSTAFF	98z	OPAL $e^+ e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.73 \pm 0.07 \pm 0.04$	<sup>9</sup> AAIJ	13K	LHCB Repl. by AAIJ 15N
$0.650 \pm 0.029 \pm 0.018$	<sup>10</sup> SAHOO	08	BELL Repl. by ADACHI 12A
$0.72 \pm 0.09 \pm 0.03$	<sup>3</sup> SAHOO	08	BELL Repl. by ADACHI 12A
$0.642 \pm 0.031 \pm 0.017$	CHEN	07	BELL Repl. by ADACHI 12A

- <sup>1</sup> AAIJ 15N uses 41,560 flavor-tagged  $B_d \rightarrow J/\psi K_S^0$  events from 3 fb<sup>-1</sup> of integrated luminosity. Provides the correlation coefficient  $\rho = 0.483$  between the statistical uncertainties of and measurements.
- <sup>2</sup> Uses  $B^0 \rightarrow J/\psi K_S^0$  decays.
- <sup>3</sup> Based on  $B^0 \rightarrow \psi(2S) K_S^0$  decays.
- <sup>4</sup> Uses  $B^0 \rightarrow J/\psi K_L^0$  decays.
- <sup>5</sup> SATO 12 uses 121 fb<sup>-1</sup> data collected at  $\Upsilon(5S)$  resonance. Uses the " $B - \pi$  tagging" where  $B\pi^+$  and  $B\pi^-$  tagged  $J/\psi K_S^0$  events are compared.
- <sup>6</sup> AFFOLDER 00c uses about 400  $B^0 \rightarrow J/\psi(1S) K_S^0$  events. The production flavor of  $B^0$  was determined using three tagging algorithms: a same-side tag, a jet-charge tag, and a soft-lepton tag.
- <sup>7</sup> BARATE 00q uses 23 candidates for  $B^0 \rightarrow J/\psi(1S) K_S^0$  decays. A combination of jet-charge, vertex-charge, and same-side tagging techniques were used to determine the  $B^0$  production flavor.
- <sup>8</sup> ACKERSTAFF 98z uses 24 candidates for  $B_d^0 \rightarrow J/\psi(1S) K_S^0$  decay. A combination of jet-charge and vertex-charge techniques were used to tag the  $B_d^0$  production flavor.
- <sup>9</sup> AAIJ 13K uses 8200 flavor-tagged  $B_d \rightarrow J/\psi K_S^0$  events from 1 fb<sup>-1</sup> of integrated luminosity. Provides the correlation coefficient  $\rho = 0.42$  between the statistical uncertainties of  $S_{J/\psi(nS) K^0}$  ( $B^0 \rightarrow J/\psi(nS) K^0$ ) and  $C_{J/\psi(nS) K^0}$  ( $B^0 \rightarrow J/\psi(nS) K^0$ ) measurements.
- <sup>10</sup> Combined result of CHEN 07 and SAHOO 08.

 $C_{J/\psi K^{*0}} (B^0 \rightarrow J/\psi K^{*0})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.025 \pm 0.083 \pm 0.054</math></b>	<sup>1</sup> AUBERT	09K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Based on $B^0 \rightarrow J/\psi K^{*0}$ , $K^{*0} \rightarrow K_S^0 \pi^0$ .			

 $S_{J/\psi K^{*0}} (B^0 \rightarrow J/\psi K^{*0})$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.601 \pm 0.239 \pm 0.087</math></b>	<sup>1,2</sup> AUBERT	09K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Based on $B^0 \rightarrow J/\psi K^{*0}$ , $K^{*0} \rightarrow K_S^0 \pi^0$ .			
<sup>2</sup> This $S_{J/\psi K^{*0}}$ value has been corrected for the dilution of the $\sin(\Delta M \Delta t)$ coefficient of the $CP$ asymmetry by a factor of $1-R_{\perp}$ , which arises from the mixture of $CP$ -even and $CP$ -odd $B$ decay amplitudes.			

 $C_{\chi_{c0} K_S^0} (B^0 \rightarrow \chi_{c0} K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.29 \pm 0.53 \pm 0.06</math></b>	<sup>1</sup> AUBERT	09AU BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

 $S_{\chi_{c0} K_S^0} (B^0 \rightarrow \chi_{c0} K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.69 \pm 0.52 \pm 0.08</math></b>	<sup>1</sup> AUBERT	09AU BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses Dalitz plot analysis of $B^0 \rightarrow K^0 \pi^+ \pi^-$ decays and the first of two equivalent solutions is used.			

 $C_{\chi_{c1} K_S^0} (B^0 \rightarrow \chi_{c1} K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.07</math> OUR AVERAGE</b>			
$0.017 \pm 0.083 \pm 0.026$	ADACHI	12A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.129 \pm 0.109 \pm 0.025$	AUBERT	09K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $S_{\chi_{c1} K_S^0} (B^0 \rightarrow \chi_{c1} K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.63 \pm 0.10</math> OUR AVERAGE</b>			
$0.640 \pm 0.117 \pm 0.040$	ADACHI	12A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.614 \pm 0.160 \pm 0.040$	AUBERT	09K BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $\sin(2\beta_{\text{eff}})(B^0 \rightarrow \phi K^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.22 \pm 0.27 \pm 0.12</math></b>	AUBERT	07AX BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.50 \pm 0.25 \pm 0.07$	<sup>1</sup> AUBERT	05T BABR	Repl. by AUBERT 07AX
<sup>1</sup> Obtained by constraining $C = 0$ .			

 $\sin(2\beta_{\text{eff}})(B^0 \rightarrow \phi K_0^*(1430)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.97 \pm 0.03</math></b>	<sup>1</sup> AUBERT	08BG BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<b><math>-0.52</math></b>			
<sup>1</sup> Measured using the $CP$ -violation phase difference $\Delta\phi_{00}$ between the $B$ and $\bar{B}$ decay amplitude.			

 $\sin(2\beta_{\text{eff}})(B^0 \rightarrow K^+ K^- K_S^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.77 \pm 0.11 \pm 0.07</math></b>	AUBERT	07AX BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
<b><math>0.04</math></b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.55 \pm 0.22 \pm 0.12$	<sup>1</sup> AUBERT	05T BABR	Repl. by AUBERT 07AX
<sup>1</sup> Obtained by constraining $C = 0$ .			

 $\sin(2\beta_{\text{eff}})(B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.45 \pm 0.28</math> OUR AVERAGE</b>			
$0.29 \pm 0.34 \pm 0.06$	AUBERT	07BH BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.78 \pm 0.44 \pm 0.22$	KROKOVNY	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$

 $2\beta_{\text{eff}}(B^0 \rightarrow J/\psi \rho^0)$ 

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>41.7 \pm 9.6 \pm 2.8</math></b>	AAIJ	15J LHCb	$pp$ at 7, 8 TeV
<b><math>-6.3</math></b>			

 $|\lambda| (B^0 \rightarrow [K_S^0 \pi^+ \pi^-]_{D^{(*)}} h^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.01 \pm 0.08 \pm 0.02</math></b>	AUBERT	07BH BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

 $|\sin(2\beta + \gamma)|$ 

$\beta$  ( $\phi_1$ ) and  $\gamma$  ( $\phi_3$ ) are angles of CKM unitarity triangle, see the review on " $CP$  Violation" in the Reviews section.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&gt;0.40</math></b>	90	<sup>1</sup> AUBERT	06V BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$>0.13$	95	<sup>2</sup> RONGA	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$>0.07$	95	<sup>2</sup> RONGA	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$>0.35$	90	<sup>3</sup> AUBERT	05Z BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$>0.69$	68	<sup>4</sup> AUBERT	04V BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$>0.58$	95	<sup>5</sup> AUBERT	04W BABR	Repl. by AUBERT 05Z

- <sup>1</sup> Uses fully reconstructed  $B^0 \rightarrow D^{(*)} \pi^+ \pi^-$  and  $D^{\pm} \rho^{\mp}$  decays and some theoretical assumptions.
- <sup>2</sup> Combines the results from fully reconstructed and partially reconstructed  $D^{(*)} \pi$  events by taking weighted averages. Assumes that systematic errors from physics parameters and fit biases in the two measurements are 100% correlated.
- <sup>3</sup> Uses partially reconstructed  $B^0 \rightarrow D^{*} \pi^+ \pi^-$  decays and some theoretical assumptions.
- <sup>4</sup> Uses fully reconstructed  $B^0 \rightarrow D^{(*)} \pi^+ \pi^-$  decays and some theoretical assumptions, such as the SU(3) symmetry relation.
- <sup>5</sup> Combining this measurement with the results from AUBERT 04V for fully reconstructed  $B^0 \rightarrow D^{(*)} \pi^+ \pi^-$  and some theoretical assumptions, such as the SU(3) symmetry relation.

 $2\beta + \gamma$ 

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>83 \pm 53 \pm 20</math></b>	<sup>1</sup> AUBERT	08AC BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Used a time-dependent Dalitz-plot analysis of  $B^0 \rightarrow D^{\mp} K^0 \pi^{\pm}$  assuming the ratio of the  $b \rightarrow u$  and  $b \rightarrow c$  decay amplitudes to be 0.3.

 $\gamma(B^0 \rightarrow D^0 K^{*0})$ 

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>162 \pm 56</math></b>	<sup>1</sup> AUBERT	09R BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

- <sup>1</sup> Uses Dalitz plot analysis of  $D^0 \rightarrow K_S^0 \pi^+ \pi^-$  decays coming from  $B^0 \rightarrow D^0 K^{*0}$  modes. The corresponding 95% CL interval is  $77^\circ < \gamma < 247^\circ$ . A 180 degree ambiguity is implied.

 $\alpha$ 

For angle  $\alpha(\phi_2)$  of the CKM unitarity triangle, see the review on " $CP$  violation" in the reviews section.

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b><math>93 \pm 5</math> OUR AVERAGE</b>			
$93.7 \pm 10.6$	<sup>1</sup> VANHOEFER	16 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
$92.4 \pm 6.0$	<sup>1</sup> AUBERT	09G BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$92.4 \pm 6.5$			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$84.9 \pm 13.5$	<sup>1</sup> VANHOEFER	14 BELL	Repl. by VANHOEFER 16
$79 \pm 7 \pm 11$	<sup>2</sup> AUBERT	10D BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$78.6 \pm 7.3$	<sup>3</sup> AUBERT	07O BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$88 \pm 17$	<sup>4</sup> SOMOV	06 BELL	Repl. by VANHOEFER 14
$100 \pm 13$	<sup>5</sup> AUBERT,B	05C BABR	Repl. by AUBERT 09G
$102 \pm 16$	<sup>6</sup> AUBERT,B	04R BABR	Repl. by AUBERT,B 05C
$-12 \pm 14$			

- <sup>1</sup> Based on an isospin analysis of the  $B \rightarrow \rho\rho$  system.
- <sup>2</sup> Obtained using the time dependent analysis of  $B^0 \rightarrow a_1(1260) \pi^+ \pi^-$  and branching fraction measurements of  $B \rightarrow a_1(1260) K$  and  $B \rightarrow K_1 \pi$ . Uses SU(3) flavor relations.
- <sup>3</sup> The angle  $\alpha_{\text{eff}}$  is obtained using the measured  $CP$  parameters of  $B^0 \rightarrow a_1(1260) \pi^+ \pi^-$  and choosing one of the four solutions that is compatible with the result of SM-based fits.
- <sup>4</sup> Obtained using isospin relation and selecting a solution closest to the CKM best fit average; the 90% CL allowed interval is  $59^\circ < \phi_2 (= \alpha) < 115^\circ$ .
- <sup>5</sup> Obtained using isospin relation and selecting a solution closest to the CKM best fit average; 90% CL allowed interval is  $79^\circ < \alpha < 123^\circ$ .
- <sup>6</sup> Obtained from the measured  $CP$  parameters of the longitudinal polarization by selecting the solution closest to the CKM best fit central value of  $\alpha = 95^\circ - 98^\circ$ .

# Meson Particle Listings

$B^0$

## T and CPT VIOLATION PARAMETERS

Measured values of the  $T$ -,  $CP$ -, and  $CP$ - $T$ -asymmetry parameters, defined as the differences in  $S_{\alpha,\beta}^{\pm}$  and  $C_{\alpha,\beta}^{\pm}$  between symmetry-transformed transitions. The indices  $\alpha = \ell^+, \ell^-$  and  $\beta = K_S^0, K_L^0$  stand for reconstructed the flavor final state and the  $CP$  final states from  $\Upsilon(4S)$  decay. The sign  $\pm$  indicates whether the decay to the flavor final state  $\alpha$  occurs before or after the decay to the  $CP$  final state.

$\Delta S_T^+(S_{\ell^-,K_S^0}^- - S_{\ell^+,K_S^0}^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-1.37 \pm 0.14 \pm 0.06$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta S_T^-(S_{\ell^-,K_S^0}^+ - S_{\ell^+,K_S^0}^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
$1.17 \pm 0.18 \pm 0.11$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta C_T^+(C_{\ell^-,K_S^0}^- - C_{\ell^+,K_S^0}^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.10 \pm 0.14 \pm 0.08$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta C_T^-(C_{\ell^-,K_S^0}^+ - C_{\ell^+,K_S^0}^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.04 \pm 0.14 \pm 0.08$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta S_{CP}^+(S_{\ell^-,K_S^0}^+ - S_{\ell^+,K_S^0}^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-1.30 \pm 0.11 \pm 0.07$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta S_{CP}^-(S_{\ell^-,K_S^0}^- - S_{\ell^+,K_S^0}^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
$1.33 \pm 0.12 \pm 0.06$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta C_{CP}^+(C_{\ell^-,K_S^0}^+ - C_{\ell^+,K_S^0}^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.07 \pm 0.09 \pm 0.03$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta C_{CP}^-(C_{\ell^-,K_S^0}^- - C_{\ell^+,K_S^0}^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.08 \pm 0.10 \pm 0.04$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta S_{CPT}^+(S_{\ell^+,K_S^0}^- - S_{\ell^+,K_S^0}^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.16 \pm 0.21 \pm 0.09$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta S_{CPT}^-(S_{\ell^+,K_S^0}^+ - S_{\ell^+,K_S^0}^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.03 \pm 0.13 \pm 0.06$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta C_{CPT}^+(C_{\ell^+,K_S^0}^- - C_{\ell^+,K_S^0}^+)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.14 \pm 0.15 \pm 0.07$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Delta C_{CPT}^-(C_{\ell^+,K_S^0}^+ - C_{\ell^+,K_S^0}^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.03 \pm 0.12 \pm 0.08$	LEES	12W	BABR $e^+e^- \rightarrow \Upsilon(4S)$

## $B^0 \rightarrow D^{*-} \ell^+ \nu_\ell$ FORM FACTORS

$R_1$  (form factor ratio  $\sim V/A_1$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.41 \pm 0.04</math> OUR AVERAGE</b>			
$1.401 \pm 0.034 \pm 0.018$	<sup>1</sup> DUNGEL	10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$1.56 \pm 0.07 \pm 0.15$	AUBERT	09A	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$1.18 \pm 0.30 \pm 0.12$	DUBOSCQ	96	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.429 \pm 0.061 \pm 0.044$	AUBERT	08R	BABR Repl. by AUBERT 09A
$1.396 \pm 0.060 \pm 0.044$	AUBERT,B	06Z	BABR Repl. by AUBERT 08R

<sup>1</sup> Uses fully reconstructed  $D^{*-} \ell^+ \nu$  events ( $\ell = e$  or  $\mu$ ).

$R_2$  (form factor ratio  $\sim A_2/A_1$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.85 \pm 0.05</math> OUR AVERAGE</b>			Error includes scale factor of 1.9.
$0.864 \pm 0.024 \pm 0.008$	<sup>1</sup> DUNGEL	10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$0.66 \pm 0.05 \pm 0.09$	AUBERT	09A	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.71 \pm 0.22 \pm 0.07$	DUBOSCQ	96	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.827 \pm 0.038 \pm 0.022$	AUBERT	08R	BABR Repl. by AUBERT 09A
$0.885 \pm 0.040 \pm 0.026$	AUBERT,B	06Z	BABR Repl. by AUBERT 08R

<sup>1</sup> Uses fully reconstructed  $D^{*-} \ell^+ \nu$  events ( $\ell = e$  or  $\mu$ ).

$\rho_{A_1}^2$  (form factor slope)

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.204 \pm 0.031</math> OUR AVERAGE</b>			
$1.214 \pm 0.034 \pm 0.009$	<sup>1</sup> DUNGEL	10	BELL $e^+e^- \rightarrow \Upsilon(4S)$
$1.22 \pm 0.02 \pm 0.07$	AUBERT	09A	BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.91 \pm 0.15 \pm 0.06$	DUBOSCQ	96	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.191 \pm 0.048 \pm 0.028$	AUBERT	08R	BABR Repl. by AUBERT 09A
$1.145 \pm 0.059 \pm 0.046$	AUBERT,B	06Z	BABR Repl. by AUBERT 08R

<sup>1</sup> Uses fully reconstructed  $D^{*-} \ell^+ \nu$  events ( $\ell = e$  or  $\mu$ ).

## PARTIAL BRANCHING FRACTIONS IN $B^0 \rightarrow K^{*0} \ell^+ \ell^-$

$B(B^0 \rightarrow K^{*0} e^+ e^-)$  ( $0.0009 < q^2 < 1.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.1 \pm 0.9 \pm 0.2 \pm 0.2</math></b>	<sup>1</sup> AAIJ	13U	LHCB $pp$ at 7 TeV

<sup>1</sup> The last uncertainty is due to uncertainties of  $B(B^0 \rightarrow J/\psi K^{*0})$  and  $B(J/\psi \rightarrow e^+ e^-)$  branching fraction measurements.

$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $0.1 < q^2 < 2.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.24 \pm 0.23 \pm 0.27</math> OUR AVERAGE</b>			Error includes scale factor of 1.6.

$1.14 \pm 0.11 \pm 0.11$	AAIJ	13Y	LHCB $pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$1.80 \pm 0.36 \pm 0.11$	AALTONEN	11AI	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.48 \pm 0.14 \pm 0.04$	<sup>1</sup> CHATRCHYAN	13BL	CMS $pp$ at 7 TeV

$1.16 \pm 0.23 \pm 0.11$	AAIJ	12U	LHCB Repl. by AAIJ 13Y
<sup>1</sup> CHATRCHYAN 13BL uses, for this bin, $1.0 < q^2 < 2.0 \text{ GeV}^2/c^4$ .			

$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $2.0 < q^2 < 4.3 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.76 \pm 0.07</math> OUR AVERAGE</b>			
$0.759 \pm 0.115 \pm 0.046$	KHACHATRY...	16D	CMS $pp$ at 8 TeV
$0.69 \pm 0.07 \pm 0.09$	AAIJ	13Y	LHCB $pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.87 \pm 0.16 \pm 0.07$	CHATRCHYAN	13BL	CMS $pp$ at 7 TeV
$0.84 \pm 0.28 \pm 0.06$	AALTONEN	11AI	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.78 \pm 0.21 \pm 0.05$	AAIJ	12U	LHCB Repl. by AAIJ 13Y

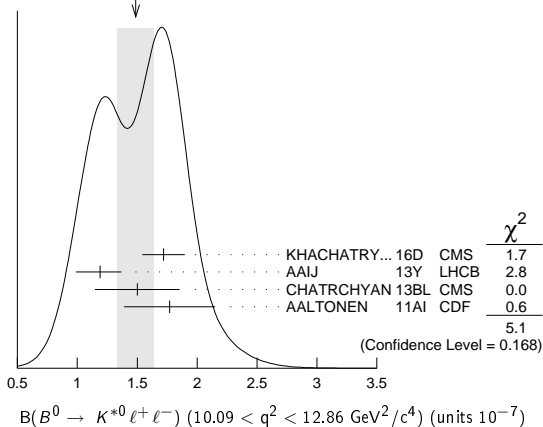
$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $4.3 < q^2 < 8.68 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.87 \pm 0.21</math> OUR AVERAGE</b>			
$2.15 \pm 0.18 \pm 0.22 \pm 0.28$	AAIJ	13Y	LHCB $pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$1.62 \pm 0.31 \pm 0.18$	CHATRCHYAN	13BL	CMS $pp$ at 7 TeV
$1.73 \pm 0.43 \pm 0.15$	AALTONEN	11AI	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.02 \pm 0.35 \pm 0.22$	AAIJ	12U	LHCB Repl. by AAIJ 13Y

$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $10.09 < q^2 < 12.86 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.49 \pm 0.15</math> OUR AVERAGE</b>			Error includes scale factor of 1.3. See the ideogram below.
$1.72 \pm 0.11 \pm 0.14$	KHACHATRY...	16D	CMS $pp$ at 8 TeV
$1.19 \pm 0.11 \pm 0.14 \pm 0.17$	AAIJ	13Y	LHCB $pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$1.50 \pm 0.25 \pm 0.25$	CHATRCHYAN	13BL	CMS $pp$ at 7 TeV
$1.77 \pm 0.36 \pm 0.12$	AALTONEN	11AI	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.52 \pm 0.25 \pm 0.19$	AAIJ	12U	LHCB Repl. by AAIJ 13Y

WEIGHTED AVERAGE  
 $1.49 \pm 0.15$  (Error scaled by 1.3)





$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $14.18 < q^2 < 16.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.09 \pm 0.10</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.		
$1.22 \pm 0.11 \pm 0.09$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$1.02 \pm 0.11 \pm_{-0.15}^{+0.11}$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.84 \pm_{-0.15}^{+0.16} \pm 0.09$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$1.34 \pm 0.26 \pm 0.08$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.15 \pm 0.20 \pm 0.09$	AAIJ	12U LHCb	Repl. by AAIJ 13Y

$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $16.0 < q^2 < 19.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.27 \pm 0.09</math> OUR AVERAGE</b>			
$1.26 \pm 0.09 \pm 0.09$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$1.23 \pm 0.12 \pm_{-0.18}^{+0.15}$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$1.56 \pm 0.18 \pm 0.15$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.97 \pm 0.26 \pm 0.07$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.50 \pm 0.24 \pm 0.15$	AAIJ	12U LHCb	Repl. by AAIJ 13Y

$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $1.0 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.78 \pm 0.15</math> OUR AVERAGE</b>			
$1.90 \pm 0.20$	KHACHATRY...16D	CMS	$pp$ at 7, 8 TeV
$1.70 \pm 0.15 \pm_{-0.25}^{+0.20}$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$1.42 \pm 0.41 \pm 0.12$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.20 \pm 0.30 \pm 0.20$	CHATRCHYAN13BL	CMS	Repl. by KHACHATRYAN 16D
$2.10 \pm 0.30 \pm 0.15$	AAIJ	12U LHCb	Repl. by AAIJ 13Y

$B(B^0 \rightarrow K^{*0} \ell^+ \ell^-)$  ( $0.0 < q^2 < 4.3 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.60 \pm 0.45 \pm 0.17</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $q^2 < 2.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.24 \pm_{-0.20}^{+0.22}</math> OUR AVERAGE</b>			
$0.21 \pm_{-0.23}^{+0.27}$	AAIJ	12AH LHCb	$pp$ at 7 TeV
$0.31 \pm 0.37 \pm 0.02$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $2.0 < q^2 < 4.3 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.24 \pm_{-0.30}^{+0.35}</math> OUR AVERAGE</b>	Error includes scale factor of 1.6.		
$0.07 \pm_{-0.21}^{+0.25}$	AAIJ	12AH LHCb	$pp$ at 7 TeV
$0.93 \pm 0.49 \pm 0.07$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $4.3 < q^2 < 8.68 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.08 \pm 0.27</math> OUR AVERAGE</b>			
$1.23 \pm 0.31$	AAIJ	12AH LHCb	$pp$ at 7 TeV
$0.66 \pm 0.51 \pm 0.05$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $10.09 < q^2 < 12.86 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.27 \pm 0.27</math> OUR AVERAGE</b>	Error includes scale factor of 1.8.		
$0.50 \pm_{-0.19}^{+0.22}$	AAIJ	12AH LHCb	$pp$ at 7 TeV
$-0.03 \pm 0.22 \pm 0.01$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $14.18 < q^2 < 16.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.29 \pm_{-0.15}^{+0.21}</math> OUR AVERAGE</b>	Error includes scale factor of 1.8.		
$0.20 \pm_{-0.09}^{+0.13}$	AAIJ	12AH LHCb	$pp$ at 7 TeV
$0.73 \pm 0.26 \pm 0.06$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $q^2 > 16.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.31 \pm_{-0.12}^{+0.16}</math> OUR AVERAGE</b>			
$0.35 \pm_{-0.14}^{+0.21}$	AAIJ	12AH LHCb	$pp$ at 7 TeV
$0.21 \pm 0.18 \pm 0.16$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $1.0 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.92 \pm 0.16</math> OUR AVERAGE</b>			
$0.916 \pm_{-0.157}^{+0.172} \pm 0.004$	<sup>1</sup> AAIJ	14M LHCb	$pp$ at 7, 8 TeV
$0.98 \pm 0.61 \pm 0.08$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.65 \pm_{-0.35}^{+0.45}$	AAIJ	12AH LHCb	Repl. by AAIJ 14M

<sup>1</sup> Uses  $B(B^0 \rightarrow J/\psi(1S) K^0) = (0.928 \pm 0.013 \pm 0.037) \times 10^{-3}$  for normalisation and  $\mu^+ \mu^-$  as a lepton pair. Measured in  $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ .

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $0.0 < q^2 < 4.3 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.27 \pm 0.62 \pm 0.10</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV

$B(B^0 \rightarrow K^0 \ell^+ \ell^-)$  ( $15.0 < q^2 < 22.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.67 \pm_{-0.11}^{+0.11} \pm 0.04</math></b>	<sup>1</sup> AAIJ	14M LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Uses  $B(B^0 \rightarrow J/\psi(1S) K^0) = (0.928 \pm 0.013 \pm 0.037) \times 10^{-3}$  for normalisation and  $\mu^+ \mu^-$  as a lepton pair.

$F_H(B^0 \rightarrow K^0 \mu^+ \mu^-)$  ( $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

$F_H$  is a fractional contribution of (pseudo) scalar and tensor amplitudes to the decay width in the massless muon approximation.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.78 \pm 0.46 \pm 0.09</math></b>	<sup>1</sup> AAIJ	14O LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> AAIJ 14O reports 68% C.L. interval, which we encode as midpoint with uncertainty as half of the width of interval.

$F_H(B^0 \rightarrow K^0 \mu^+ \mu^-)$  ( $15.0 < q^2 < 22.0 \text{ GeV}^2/c^4$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.34 \pm 0.25 \pm 0.03</math></b>	<sup>1</sup> AAIJ	14O LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> AAIJ 14O reports 68% C.L. interval, which we encode as midpoint with uncertainty as half of the width of interval.

PRODUCTION ASYMMETRIES

$A_P(B^0)$

$A_P(B^0) = [\sigma(\bar{B}^0) - \sigma(B^0)] / [\sigma(\bar{B}^0) + \sigma(B^0)]$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>-0.35 \pm 0.76 \pm 0.28</math></b>	<sup>1</sup> AAIJ	14BP LHCb	$pp$ at 7 TeV

<sup>1</sup> Based on time-dependent analysis of  $B^0 \rightarrow J/\psi K^{*0}$  and  $B^0 \rightarrow D^- \pi^+$  in kinematic range  $4 < p_T < 30 \text{ GeV}/c$  and  $2.5 < \eta < 4.5$ .

$B^0$  REFERENCES

AAIJ	16	JHEP 1601 012	R. Aaij <i>et al.</i>	(LHCb Collab.)
CHOBANOVA	16	PR D93 031101	V. Chobanova <i>et al.</i>	(BELLE Collab.)
GLATTAUER	16	PR D93 032006	R. Glattauer <i>et al.</i>	(BELLE Collab.)
HAMER	16	PR D93 032007	P. Hamer <i>et al.</i>	(BELLE Collab.)
KHACHATRY...	16D	PL B753 424	V. Khachatryan <i>et al.</i>	(CMS Collab.)
LEES	16	PRL 116 041801	J.P. Lees <i>et al.</i>	(BABAR Collab.)
VANHOFER	16	PR D93 032010	P. Vanhoefer <i>et al.</i>	(BELLE Collab.)
AAIJ	15AC	JHEP 1505 019	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15AS	JHEP 1510 053	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15AZ	PRL 115 161802	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15BB	PR D92 112002	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15D	JHEP 1501 024	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15F	PRL 114 041601	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15J	PL B742 38	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15N	PRL 115 031601	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15Q	PRL 115 111803	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15S	PL B743 46	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15T	PL B747 468	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15X	PR D92 012012	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15Y	PR D92 032002	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15Z	JHEP 1504 064	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	15A	PRL 114 062001	V.M. Abazov <i>et al.</i>	(DO Collab.)
ABDESSALAM	15	PRL 115 121604	A. Abdessalam <i>et al.</i>	(BABAR and BELLE Collabs.)
BALA	15	PR D91 051101	A. Bala <i>et al.</i>	(BELLE Collab.)







See key on page 601

# Meson Particle Listings

## $B^0, B^\pm/B^0$ ADMIXTURE

BEAN	93B	PRL 70 2681	A. Bean <i>et al.</i>	(CLEO Collab.)
BUSKULIC	93D	PL B307 194	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
Also		PL B325 537 (erratum)	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	93K	PL B313 498	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
SANGHERA	93	PR D47 791	S. Sanghera <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92C	PL B275 195	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92G	ZPHY C54 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92L	ZPHY C55 357	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BORTOLETTO	92	PR D45 21	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
HENDERSON	92	PR D45 2212	S. Henderson <i>et al.</i>	(CLEO Collab.)
KRAMER	92	PL B279 181	G. Kramer, W.F. Palmer	(HAMB, OSU)
ALBAJAR	91C	PL B262 163	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALBAJAR	91E	PL B273 540	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALBRECHT	91B	PL B254 288	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	91C	PL B255 297	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	91E	PL B262 148	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BERKELMAN	91	ARNPS 41 1	K. Berkelman, S. Stone	(CORN, SYRA)
"Decays of $B$ Mesons"				
FULTON	91	PR D43 651	R. Fulton <i>et al.</i>	(CLEO Collab.)
ALBRECHT	90B	PL B241 278	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	90J	ZPHY C48 543	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANTREASANYAN	90B	ZPHY C48 553	D. Antreasanyan <i>et al.</i>	(Crystal Ball Collab.)
BORTOLETTO	90	PRL 64 2117	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
ELSEN	90	ZPHY C46 349	E. Elsen <i>et al.</i>	(JADE Collab.)
ROSNER	90	PR D42 3732	J.L. Rosner	
WAGNER	90	PRL 64 1095	S.R. Wagner <i>et al.</i>	(Mark II Collab.)
ALBRECHT	89C	PL B219 121	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	89G	PL B229 304	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	89J	PL B229 175	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	89L	PL B232 554	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ARTUSO	89	PRL 62 2233	M. Artuso <i>et al.</i>	(CLEO Collab.)
AVERILL	89	PR D39 123	D.A. Averill <i>et al.</i>	(HRS Collab.)
AVERY	89B	PL B223 470	P. Avery <i>et al.</i>	(CLEO Collab.)
BEBEK	89	PRL 62 8	C. Bebek <i>et al.</i>	(CLEO Collab.)
BORTOLETTO	89	PRL 62 2436	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
BORTOLETTO	89B	PRL 63 1667	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
ALBRECHT	88F	PL B209 119	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88K	PL B215 424	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87C	PL B185 218	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87D	PL B199 451	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87I	PL B192 245	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87J	PL B197 452	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AVERY	87	PL B183 429	P. Avery <i>et al.</i>	(CLEO Collab.)
BEAN	87B	PRL 58 183	A. Bean <i>et al.</i>	(CLEO Collab.)
BEBEK	87	PR D36 1289	C. Bebek <i>et al.</i>	(CLEO Collab.)
ALAM	86	PR D34 3279	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	86F	PL B182 95	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
PDG	86	PL 170B 1	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)
CHEN	85	PR D31 2386	A. Chen <i>et al.</i>	(CLEO Collab.)
HAAS	85	PRL 55 1248	J. Haas <i>et al.</i>	(CLEO Collab.)
AVERY	84	PRL 53 1309	P. Avery <i>et al.</i>	(CLEO Collab.)
GILES	84	PR D30 2279	R. Giles <i>et al.</i>	(CLEO Collab.)
BEHRENDTS	83	PRL 50 681	S. Behrendts <i>et al.</i>	(CLEO Collab.)

## $B^\pm/B^0$ ADMIXTURE

### B DECAY MODES

The branching fraction measurements are for an admixture of  $B$  mesons that the  $\Upsilon(4S)$ . The values quoted assume that  $\mathcal{B}(\Upsilon(4S) \rightarrow B\bar{B}) = 100\%$ .

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm \text{ anything}$ , the treatment of multiple  $D$ 's in the final state must be defined. One possibility would be to count the number of events with one-or-more  $D$ 's and divide by the total number of  $B$ 's. Another possibility would be to count the total number of  $D$ 's and divide by the total number of  $B$ 's, which is the definition of average multiplicity. The two definitions are identical if only one  $D$  is allowed in the final state. Even though the "one-or-more" definition seems sensible, for practical reasons inclusive branching fractions are almost always measured using the multiplicity definition. For heavy final state particles, authors call their results inclusive branching fractions while for light particles some authors call their results multiplicities. In the  $B$  sections, we list all results as inclusive branching fractions, adopting a multiplicity definition. This means that inclusive branching fractions can exceed 100% and that inclusive partial widths can exceed total widths, just as inclusive cross sections can exceed total cross section.

$\bar{B}$  modes are charge conjugates of the modes below. Reactions indicate the weak decay vertex and do not include mixing.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Semileptonic and leptonic modes</b>		
$\Gamma_1$ $e^+ \nu_e$ anything	[a]	
$\Gamma_2$ $\mu^+ \nu_\mu$ anything	[a]	
$\Gamma_3$ $\ell^+ \nu_\ell$ anything	[a,b] ( 10.86 $\pm$ 0.16 ) %	
$\Gamma_4$ $D^- \ell^+ \nu_\ell$ anything	[b] ( 2.8 $\pm$ 0.9 ) %	
$\Gamma_5$ $\bar{D}^0 \ell^+ \nu_\ell$ anything	[b] ( 7.3 $\pm$ 1.5 ) %	
$\Gamma_6$ $\bar{D} \ell^+ \nu_\ell$	( 2.42 $\pm$ 0.12 ) %	
$\Gamma_7$ $D^{*-} \ell^+ \nu_\ell$ anything	[c] ( 6.7 $\pm$ 1.3 ) $\times 10^{-3}$	
$\Gamma_8$ $D^{*0} \ell^+ \nu_\ell$ anything		
$\Gamma_9$ $D^* \ell^+ \nu_\ell$	[d] ( 4.95 $\pm$ 0.11 ) %	
$\Gamma_{10}$ $\bar{D}^{**} \ell^+ \nu_\ell$	[b,e] ( 2.7 $\pm$ 0.7 ) %	
$\Gamma_{11}$ $D_1(2420) \ell^+ \nu_\ell$ anything	( 3.8 $\pm$ 1.3 ) $\times 10^{-3}$	S=2.4
$\Gamma_{12}$ $D \pi \ell^+ \nu_\ell$ anything + $D^* \pi \ell^+ \nu_\ell$ anything	( 2.6 $\pm$ 0.5 ) %	S=1.5
$\Gamma_{13}$ $D \pi \ell^+ \nu_\ell$ anything	( 1.5 $\pm$ 0.6 ) %	
$\Gamma_{14}$ $D^* \pi \ell^+ \nu_\ell$ anything	( 1.9 $\pm$ 0.4 ) %	

$\Gamma_{15}$ $\bar{D}_2^*(2460) \ell^+ \nu_\ell$ anything	( 4.4 $\pm$ 1.6 ) $\times 10^{-3}$	
$\Gamma_{16}$ $D^{*-} \pi^+ \ell^+ \nu_\ell$ anything	( 1.00 $\pm$ 0.34 ) %	
$\Gamma_{17}$ $\bar{D} \pi^+ \pi^- \ell^+ \nu_\ell$	( 1.62 $\pm$ 0.32 ) $\times 10^{-3}$	
$\Gamma_{18}$ $\bar{D}^* \pi^+ \pi^- \ell^+ \nu_\ell$	( 9.4 $\pm$ 3.2 ) $\times 10^{-4}$	
$\Gamma_{19}$ $D_s^- \ell^+ \nu_\ell$ anything	[b] < 7 $\times 10^{-3}$	CL=90%
$\Gamma_{20}$ $D_s^- \ell^+ \nu_\ell K^+$ anything	[b] < 5 $\times 10^{-3}$	CL=90%
$\Gamma_{21}$ $D_s^- \ell^+ \nu_\ell K^0$ anything	[b] < 7 $\times 10^{-3}$	CL=90%
$\Gamma_{22}$ $X_c \ell^+ \nu_\ell$	( 10.65 $\pm$ 0.16 ) %	
$\Gamma_{23}$ $X_u \ell^+ \nu_\ell$	( 2.14 $\pm$ 0.31 ) $\times 10^{-3}$	
$\Gamma_{24}$ $K^+ \ell^+ \nu_\ell$ anything	[b] ( 6.3 $\pm$ 0.6 ) %	
$\Gamma_{25}$ $K^- \ell^+ \nu_\ell$ anything	[b] ( 10 $\pm$ 4 ) $\times 10^{-3}$	
$\Gamma_{26}$ $K^0/\bar{K}^0 \ell^+ \nu_\ell$ anything	[b] ( 4.6 $\pm$ 0.5 ) %	
$\Gamma_{27}$ $\bar{D} \tau^+ \nu_\tau$	( 9.8 $\pm$ 1.3 ) $\times 10^{-3}$	
$\Gamma_{28}$ $D^* \tau^+ \nu_\tau$	( 1.58 $\pm$ 0.12 ) %	

### D, D\*, or D<sub>s</sub> modes

$\Gamma_{29}$ $D^\pm$ anything	( 22.9 $\pm$ 1.3 ) %	
$\Gamma_{30}$ $D^0/\bar{D}^0$ anything	( 61.8 $\pm$ 2.9 ) %	S=1.3
$\Gamma_{31}$ $D^*(2010)^\pm$ anything	( 22.5 $\pm$ 1.5 ) %	
$\Gamma_{32}$ $D^*(2007)^0$ anything	( 26.0 $\pm$ 2.7 ) %	
$\Gamma_{33}$ $D_s^\pm$ anything	[f] ( 8.3 $\pm$ 0.8 ) %	
$\Gamma_{34}$ $D_s^{*\pm}$ anything	( 6.3 $\pm$ 1.0 ) %	
$\Gamma_{35}$ $D_s^\pm \bar{D}^{(*)}$	( 3.4 $\pm$ 0.6 ) %	
$\Gamma_{36}$ $\bar{D} D_{s0}(2317)$	seen	
$\Gamma_{37}$ $\bar{D} D_{sJ}(2457)$	seen	
$\Gamma_{38}$ $D^{(*)} \bar{D}^{(*)} K^0 + D^{(*)} \bar{D}^{(*)} K^\pm [f_g]$	( 7.1 $\pm$ 2.7 ) % ( 1.7 )	
$\Gamma_{39}$ $b \rightarrow c \bar{c} s$	( 22 $\pm$ 4 ) %	
$\Gamma_{40}$ $D_s^{(*)} \bar{D}^{(*)}$	[f <sub>g</sub> ] ( 3.9 $\pm$ 0.4 ) %	
$\Gamma_{41}$ $D^* D^*(2010)^\pm$	[f] < 5.9 $\times 10^{-3}$	CL=90%
$\Gamma_{42}$ $D^* D^*(2010)^\pm + D^* D^\pm$	[f] < 5.5 $\times 10^{-3}$	CL=90%
$\Gamma_{43}$ $D D^\pm$	[f] < 3.1 $\times 10^{-3}$	CL=90%
$\Gamma_{44}$ $D_s^{(*)} \pm \bar{D}^{(*)} X (n \pi^\pm)$	[f <sub>g</sub> ] ( 9 $\pm$ 5 ) % ( 4 )	
$\Gamma_{45}$ $D^*(2010) \gamma$	< 1.1 $\times 10^{-3}$	CL=90%
$\Gamma_{46}$ $D_s^+ \pi^-, D_s^{*+} \pi^-, D_s^+ \rho^-, D_s^{*+} \rho^-, D_s^+ \pi^0, D_s^{*+} \pi^0, D_s^+ \eta, D_s^{*+} \eta, D_s^+ \rho^0, D_s^{*+} \rho^0, D_s^+ \omega, D_s^{*+} \omega$	[f] < 4 $\times 10^{-4}$	CL=90%
$\Gamma_{47}$ $D_{s1}(2536)^+ \text{ anything}$	< 9.5 $\times 10^{-3}$	CL=90%

### Charmonium modes

$\Gamma_{48}$ $J/\psi(1S)$ anything	( 1.094 $\pm$ 0.032 ) %	S=1.1
$\Gamma_{49}$ $J/\psi(1S)$ (direct) anything	( 7.8 $\pm$ 0.4 ) $\times 10^{-3}$	S=1.1
$\Gamma_{50}$ $\psi(2S)$ anything	( 3.07 $\pm$ 0.21 ) $\times 10^{-3}$	
$\Gamma_{51}$ $\chi_{c1}(1P)$ anything	( 3.86 $\pm$ 0.27 ) $\times 10^{-3}$	
$\Gamma_{52}$ $\chi_{c1}(1P)$ (direct) anything	( 3.24 $\pm$ 0.25 ) $\times 10^{-3}$	
$\Gamma_{53}$ $\chi_{c2}(1P)$ anything	( 1.4 $\pm$ 0.4 ) $\times 10^{-3}$	S=1.9
$\Gamma_{54}$ $\chi_{c2}(1P)$ (direct) anything	( 1.65 $\pm$ 0.31 ) $\times 10^{-3}$	
$\Gamma_{55}$ $\eta_c(1S)$ anything	< 9 $\times 10^{-3}$	CL=90%
$\Gamma_{56}$ $KX(3872), X \rightarrow D^0 \bar{D}^0 \pi^0$	( 1.2 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{57}$ $KX(3872), X \rightarrow D^{*0} \bar{D}^0$	( 8.0 $\pm$ 2.2 ) $\times 10^{-5}$	
$\Gamma_{58}$ $KX(3940), X \rightarrow D^{*0} \bar{D}^0$	< 6.7 $\times 10^{-5}$	CL=90%
$\Gamma_{59}$ $KX(3915), X \rightarrow \omega J/\psi$	[h] ( 7.1 $\pm$ 3.4 ) $\times 10^{-5}$	

### K or K\* modes

$\Gamma_{60}$ $K^\pm$ anything	[f] ( 78.9 $\pm$ 2.5 ) %	
$\Gamma_{61}$ $K^+$ anything	( 66 $\pm$ 5 ) %	
$\Gamma_{62}$ $K^-$ anything	( 13 $\pm$ 4 ) %	
$\Gamma_{63}$ $K^0/\bar{K}^0$ anything	[f] ( 64 $\pm$ 4 ) %	
$\Gamma_{64}$ $K^*(892)^\pm$ anything	( 18 $\pm$ 6 ) %	
$\Gamma_{65}$ $K^*(892)^0/\bar{K}^*(892)^0$ anything	[f] ( 14.6 $\pm$ 2.6 ) %	
$\Gamma_{66}$ $K^*(892) \gamma$	( 4.2 $\pm$ 0.6 ) $\times 10^{-5}$	
$\Gamma_{67}$ $\eta K \gamma$	( 8.5 $\pm$ 1.8 ) $\times 10^{-6}$ ( 1.6 )	
$\Gamma_{68}$ $K_1(1400) \gamma$	< 1.27 $\times 10^{-4}$	CL=90%
$\Gamma_{69}$ $K_2^*(1430) \gamma$	( 1.7 $\pm$ 0.6 ) $\times 10^{-5}$ ( 0.5 )	
$\Gamma_{70}$ $K_2(1770) \gamma$	< 1.2 $\times 10^{-3}$	CL=90%
$\Gamma_{71}$ $K_3^*(1780) \gamma$	< 3.7 $\times 10^{-5}$	CL=90%
$\Gamma_{72}$ $K_4^*(2045) \gamma$	< 1.0 $\times 10^{-3}$	CL=90%
$\Gamma_{73}$ $K \eta(958)$	( 8.3 $\pm$ 1.1 ) $\times 10^{-5}$	
$\Gamma_{74}$ $K^*(892) \eta/(958)$	( 4.1 $\pm$ 1.1 ) $\times 10^{-6}$	
$\Gamma_{75}$ $K \eta$	< 5.2 $\times 10^{-6}$	CL=90%
$\Gamma_{76}$ $K^*(892) \eta$	( 1.8 $\pm$ 0.5 ) $\times 10^{-5}$	

## Meson Particle Listings

 $B^\pm/B^0$  ADMIXTURE

$\Gamma_{77}$	$K\phi\phi$	(	$2.3 \pm 0.9$	) $\times 10^{-6}$	
$\Gamma_{78}$	$\bar{b} \rightarrow \bar{s}\gamma$	(	$3.49 \pm 0.19$	) $\times 10^{-4}$	
$\Gamma_{79}$	$\bar{b} \rightarrow \bar{d}\gamma$	(	$9.2 \pm 3.0$	) $\times 10^{-6}$	
$\Gamma_{80}$	$\bar{b} \rightarrow \bar{s}\text{gluon}$	<	6.8	%	CL=90%
$\Gamma_{81}$	$\eta$ anything	(	$2.6 \pm 0.5$	) $\times 10^{-4}$	
$\Gamma_{82}$	$\eta'$ anything	(	$4.2 \pm 0.9$	) $\times 10^{-4}$	
$\Gamma_{83}$	$K^+\text{gluon (charmless)}$	<	1.87	$\times 10^{-4}$	CL=90%
$\Gamma_{84}$	$K^0\text{gluon (charmless)}$	(	$1.9 \pm 0.7$	) $\times 10^{-4}$	

## Light unflavored meson modes

$\Gamma_{85}$	$\rho\gamma$	(	$1.39 \pm 0.25$	) $\times 10^{-6}$	S=1.2
$\Gamma_{86}$	$\rho/\omega\gamma$	(	$1.30 \pm 0.23$	) $\times 10^{-6}$	S=1.2
$\Gamma_{87}$	$\pi^\pm$ anything	[f] (	$358 \pm 7$	) %	
$\Gamma_{88}$	$\pi^0$ anything	(	$235 \pm 11$	) %	
$\Gamma_{89}$	$\eta$ anything	(	$17.6 \pm 1.6$	) %	
$\Gamma_{90}$	$\rho^0$ anything	(	$21 \pm 5$	) %	
$\Gamma_{91}$	$\omega$ anything	<	81	%	CL=90%
$\Gamma_{92}$	$\phi$ anything	(	$3.43 \pm 0.12$	) %	
$\Gamma_{93}$	$\phi K^*(892)$	<	2.2	$\times 10^{-5}$	CL=90%
$\Gamma_{94}$	$\bar{b} \rightarrow \bar{d}\text{gluon}$				
$\Gamma_{95}$	$\pi^+\text{gluon (charmless)}$	(	$3.7 \pm 0.8$	) $\times 10^{-4}$	

## Baryon modes

$\Gamma_{96}$	$\Lambda_c^+ / \bar{\Lambda}_c^-$ anything	(	$3.5 \pm 0.4$	) %	
$\Gamma_{97}$	$\Lambda_c^+$ anything	<	1.3	%	CL=90%
$\Gamma_{98}$	$\bar{\Lambda}_c^-$ anything	<	7	%	CL=90%
$\Gamma_{99}$	$\bar{\Lambda}_c^- \ell^+$ anything	<	9	$\times 10^{-4}$	CL=90%
$\Gamma_{100}$	$\bar{\Lambda}_c^- e^+$ anything	<	1.8	$\times 10^{-3}$	CL=90%
$\Gamma_{101}$	$\bar{\Lambda}_c^- \mu^+$ anything	<	1.4	$\times 10^{-3}$	CL=90%
$\Gamma_{102}$	$\bar{\Lambda}_c^- p$ anything	(	$2.02 \pm 0.33$	) %	
$\Gamma_{103}$	$\bar{\Lambda}_c^- p e^+ \nu_e$	<	8	$\times 10^{-4}$	CL=90%
$\Gamma_{104}$	$\Sigma_c^{++}$ anything	(	$3.3 \pm 1.7$	) $\times 10^{-3}$	
$\Gamma_{105}$	$\bar{\Sigma}_c^{--}$ anything	<	8	$\times 10^{-3}$	CL=90%
$\Gamma_{106}$	$\bar{\Sigma}_c^{--}$ anything	(	$3.6 \pm 1.7$	) $\times 10^{-3}$	
$\Gamma_{107}$	$\bar{\Sigma}_c^{--} N (N = p \text{ or } n)$	<	1.2	$\times 10^{-3}$	CL=90%
$\Gamma_{108}$	$\Xi_c^0$ anything, $\Xi_c^0 \rightarrow \Xi^- \pi^+$	(	$1.93 \pm 0.30$	) $\times 10^{-4}$	S=1.1
$\Gamma_{109}$	$\Xi_c^+, \Xi_c^+ \rightarrow \Xi^- \pi^+ \pi^+$	(	$4.5 \pm 1.3$	) $\times 10^{-4}$	
$\Gamma_{110}$	$p/\bar{p}$ anything	[f] (	$8.0 \pm 0.4$	) %	
$\Gamma_{111}$	$p/\bar{p}$ (direct) anything	[f] (	$5.5 \pm 0.5$	) %	
$\Gamma_{112}$	$\bar{p} e^+ \nu_e$ anything	<	5.9	$\times 10^{-4}$	CL=90%
$\Gamma_{113}$	$\Lambda/\bar{\Lambda}$ anything	[f] (	$4.0 \pm 0.5$	) %	
$\Gamma_{114}$	$\Lambda$ anything	seen			
$\Gamma_{115}$	$\bar{\Lambda}$ anything	seen			
$\Gamma_{116}$	$\Xi^- / \Xi^+$ anything	[f] (	$2.7 \pm 0.6$	) $\times 10^{-3}$	
$\Gamma_{117}$	baryons anything	(	$6.8 \pm 0.6$	) %	
$\Gamma_{118}$	$p\bar{p}$ anything	(	$2.47 \pm 0.23$	) %	
$\Gamma_{119}$	$\Lambda\bar{p}/\bar{\Lambda}p$ anything	[f] (	$2.5 \pm 0.4$	) %	
$\Gamma_{120}$	$\Lambda\bar{\Lambda}$ anything	<	5	$\times 10^{-3}$	CL=90%

Lepton Family number (LF) violating modes or  $\Delta B = 1$  weak neutral current ( $B1$ ) modes

$\Gamma_{121}$	$s e^+ e^-$	$B1$	(	$6.7 \pm 1.7$	) $\times 10^{-6}$	S=2.0
$\Gamma_{122}$	$s \mu^+ \mu^-$	$B1$	(	$4.3 \pm 1.0$	) $\times 10^{-6}$	
$\Gamma_{123}$	$s \ell^+ \ell^-$	$B1$	[b] (	$5.8 \pm 1.3$	) $\times 10^{-6}$	S=1.8
$\Gamma_{124}$	$\pi \ell^+ \ell^-$	$B1$	<	5.9	$\times 10^{-8}$	CL=90%
$\Gamma_{125}$	$\pi e^+ e^-$	$B1$	<	1.10	$\times 10^{-7}$	CL=90%
$\Gamma_{126}$	$\pi \mu^+ \mu^-$	$B1$	<	5.0	$\times 10^{-8}$	CL=90%
$\Gamma_{127}$	$K e^+ e^-$	$B1$	(	$4.4 \pm 0.6$	) $\times 10^{-7}$	
$\Gamma_{128}$	$K^*(892) e^+ e^-$	$B1$	(	$1.19 \pm 0.20$	) $\times 10^{-6}$	S=1.2
$\Gamma_{129}$	$K \mu^+ \mu^-$	$B1$	(	$4.4 \pm 0.4$	) $\times 10^{-7}$	
$\Gamma_{130}$	$K^*(892) \mu^+ \mu^-$	$B1$	(	$1.06 \pm 0.09$	) $\times 10^{-6}$	
$\Gamma_{131}$	$K \ell^+ \ell^-$	$B1$	(	$4.8 \pm 0.4$	) $\times 10^{-7}$	
$\Gamma_{132}$	$K^*(892) \ell^+ \ell^-$	$B1$	(	$1.05 \pm 0.10$	) $\times 10^{-6}$	
$\Gamma_{133}$	$K \nu \bar{\nu}$	$B1$	<	1.7	$\times 10^{-5}$	CL=90%
$\Gamma_{134}$	$K^* \nu \bar{\nu}$	$B1$	<	7.6	$\times 10^{-5}$	CL=90%
$\Gamma_{135}$	$s e^\pm \mu^\mp$	LF	[f] <	2.2	$\times 10^{-5}$	CL=90%
$\Gamma_{136}$	$\pi e^\pm \mu^\mp$	LF	<	9.2	$\times 10^{-8}$	CL=90%
$\Gamma_{137}$	$\rho e^\pm \mu^\mp$	LF	<	3.2	$\times 10^{-6}$	CL=90%
$\Gamma_{138}$	$K e^\pm \mu^\mp$	LF	<	3.8	$\times 10^{-8}$	CL=90%
$\Gamma_{139}$	$K^*(892) e^\pm \mu^\mp$	LF	<	5.1	$\times 10^{-7}$	CL=90%

[a] These values are model dependent.

[b] An  $\ell$  indicates an  $e$  or a  $\mu$  mode, not a sum over these modes.

[c] Here "anything" means at least one particle observed.

[d] This is a  $B(B^0 \rightarrow D^{*-} \ell^+ \nu_\ell)$  value.[e]  $D^{**}$  stands for the sum of the  $D(1^1P_1)$ ,  $D(1^3P_0)$ ,  $D(1^3P_1)$ ,  $D(1^3P_2)$ ,  $D(2^1S_0)$ , and  $D(2^1S_1)$  resonances.

[f] The value is for the sum of the charge states or particle/antiparticle states indicated.

[g]  $D^{(*)}\bar{D}^{(*)}$  stands for the sum of  $D^*\bar{D}^*$ ,  $D^*\bar{D}$ ,  $D\bar{D}^*$ , and  $D\bar{D}$ .[h]  $X(3915)$  denotes a near-threshold enhancement in the  $\omega J/\psi$  mass spectrum.

[i] Inclusive branching fractions have a multiplicity definition and can be greater than 100%.

 $B^\pm/B^0$  ADMIXTURE BRANCHING RATIOS $\Gamma(\ell^+ \nu_\ell \text{ anything})/\Gamma_{\text{total}}$  $\Gamma_3/\Gamma$ 

These branching fractions are model dependent.

"OUR EVALUATION" assumes lepton universality and is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.1086 ± 0.0016 OUR EVALUATION</b>			
<b>0.1044 ± 0.0025 OUR AVERAGE</b>			Error includes scale factor of 1.5. See the ideogram below.
0.1028 ± 0.0018 ± 0.0024	1	URQUIJO	07 BELL $e^+ e^- \rightarrow \gamma(4S)$
0.0996 ± 0.0019 ± 0.0032	2	AUBERT,B	06Y BABR $e^+ e^- \rightarrow \gamma(4S)$
0.1091 ± 0.0009 ± 0.0024	3	MAHMOOD	04 CLEO $e^+ e^- \rightarrow \gamma(4S)$
0.097 ± 0.005 ± 0.004	4	ALBRECHT	93H ARG $e^+ e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1085 ± 0.0021 ± 0.0036	5	OKABE	05 BELL Repl. by URQUIJO 07
0.1083 ± 0.0016 ± 0.0006	6	AUBERT	04X BABR Repl. by AUBERT,B 06Y
0.1036 ± 0.0006 ± 0.0023	7	AUBERT,B	04A BABR $e^+ e^- \rightarrow \gamma(4S)$
0.1087 ± 0.0018 ± 0.0030	8	AUBERT	03 BABR Repl. by AUBERT 04X
0.109 ± 0.0012 ± 0.0049	9	ABE	02Y BELL Repl. by OKABE 05
0.1049 ± 0.0017 ± 0.0043	10	BARISH	96B CLE2 Repl. by MAHMOOD 04
0.108 ± 0.002 ± 0.0056	11	HENDERSON	92 CLEO $e^+ e^- \rightarrow \gamma(4S)$
0.100 ± 0.004 ± 0.003	12	YANAGISAWA	91 CSB2 $e^+ e^- \rightarrow \gamma(4S)$
0.103 ± 0.006 ± 0.002	13	ALBRECHT	90H ARG Direct $e$ at $\gamma(4S)$
0.100 ± 0.006 ± 0.002	14	ALBRECHT	90H ARG Direct $\mu$ at $\gamma(4S)$
0.117 ± 0.004 ± 0.010	15	WACHS	89 CBAL Direct $e$ at $\gamma(4S)$
0.120 ± 0.007 ± 0.005		CHEN	84 CLEO Direct $e$ at $\gamma(4S)$
0.108 ± 0.006 ± 0.01		CHEN	84 CLEO Direct $\mu$ at $\gamma(4S)$
0.112 ± 0.009 ± 0.01		LEVMAN	84 CUSB Direct $\mu$ at $\gamma(4S)$
0.132 ± 0.008 ± 0.014	16	KLOPFEN...	83B CUSB Direct $e$ at $\gamma(4S)$

1 URQUIJO 07 report a measurement of  $(10.07 \pm 0.18 \pm 0.21)\%$  for the partial branching fraction of  $B \rightarrow e \nu_e X_c$  decay with electron energy above 0.6 GeV. We converted the result to  $B \rightarrow e \nu_e X$  branching fraction.

2 The measurements are obtained for charged and neutral  $B$  mesons partial rates of semi-leptonic decay to electrons with momentum above 0.6 GeV/c in the  $B$  rest frame. The best precision on the ratio is achieved for a momentum threshold of 1.0 GeV:  $B(B^+ \rightarrow e^+ \nu_e X) / B(B^0 \rightarrow e^+ \nu_e X) = 1.074 \pm 0.041 \pm 0.026$ .

3 Uses charge and angular correlations in  $\gamma(4S)$  events with a high-momentum lepton and an additional electron.

4 ALBRECHT 93H analysis performed using tagged semileptonic decays of the  $B$ . This technique is almost model independent for the lepton branching ratio.

5 The measurements are obtained for charged and neutral  $B$  mesons partial rates of semi-leptonic decay to electrons with momentum above 0.6 GeV/c in the  $B$  rest frame, and their ratio of  $B(B^+ \rightarrow e^+ \nu_e X)/B(B^0 \rightarrow e^+ \nu_e X) = 1.08 \pm 0.05 \pm 0.02$ .

6 The semileptonic branching ratio,  $|V_{cb}|$  and other heavy-quark parameters are determined from a simultaneous fit to moments of the hadronic-mass and lepton-energy distribution.

7 Uses the high-momentum lepton tag method and requires the electron energy above 0.6 GeV.

8 Uses the high-momentum lepton tag method. They also report  $|V_{cb}| = 0.0423 \pm 0.0007(\text{exp}) \pm 0.0020(\text{theo.})$ .

9 Uses the high-momentum lepton tag method. ABE 02Y also reports  $|V_{cb}| = 0.0408 \pm 0.0010(\text{exp}) \pm 0.0025(\text{theo.})$ . The second error is due to uncertainties of theoretical inputs.

10 BARISH 96B analysis performed using tagged semileptonic decays of the  $B$ . This technique is almost model independent for the lepton branching ratio.

11 HENDERSON 92 measurement employs  $e$  and  $\mu$ . The systematic error contains 0.004 in quadrature from model dependence. The authors average a variation of the Isgur, Scora, Grinstein, and Wise model with that of the Altarelli-Cabibbo-Corbo-Maiani-Martinelli model for semileptonic decays to correct the acceptance.

12 YANAGISAWA 91 also measures an average semileptonic branching ratio at the  $\gamma(5S)$  of 9.6–10.5% depending on assumptions about the relative production of different  $B$  meson species.

13 ALBRECHT 90H uses the model of ALTARELLI 82 to correct over all lepton momenta.  $0.099 \pm 0.006$  is obtained using ISGUR 89b.

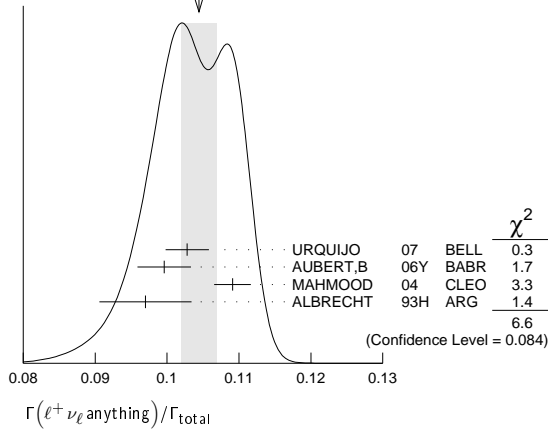
14 ALBRECHT 90H uses the model of ALTARELLI 82 to correct over all lepton momenta.  $0.097 \pm 0.006$  is obtained using ISGUR 89b.

15 Using data above  $p(e) = 2.4$  GeV, WACHS 89 determine  $\sigma(B \rightarrow e \nu \mu p)/\sigma(B \rightarrow e \nu \text{charm}) < 0.065$  at 90% CL.

16 Ratio  $\sigma(B \rightarrow e \nu \mu p)/\sigma(B \rightarrow e \nu \text{charm}) < 0.055$  at CL = 90%.

See key on page 601

## Meson Particle Listings

 $B^\pm/B^0$  ADMIXTUREWEIGHTED AVERAGE  
0.1044±0.0025 (Error scaled by 1.5) $\Gamma(D^-\ell^+\nu_\ell \text{ anything})/\Gamma(\ell^+\nu_\ell \text{ anything})$   $\Gamma_4/\Gamma_3$   
 $\ell = e \text{ or } \mu$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.26±0.07±0.04</b>	<sup>1</sup> FULTON	91	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>FULTON 91 uses  $B(D^+ \rightarrow K^-\pi^+\pi^+) = (9.1 \pm 1.3 \pm 0.4)\%$  as measured by MARK III. $\Gamma(\bar{D}^0\ell^+\nu_\ell \text{ anything})/\Gamma(\ell^+\nu_\ell \text{ anything})$   $\Gamma_5/\Gamma_3$   
 $\ell = e \text{ or } \mu$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.67±0.09±0.10</b>	<sup>1</sup> FULTON	91	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>FULTON 91 uses  $B(D^0 \rightarrow K^-\pi^+) = (4.2 \pm 0.4 \pm 0.4)\%$  as measured by MARK III. $\Gamma(\bar{D}\ell^+\nu_\ell)/\Gamma(\ell^+\nu_\ell \text{ anything})$   $\Gamma_6/\Gamma_3$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.223±0.006±0.009</b>	<sup>1</sup> AUBERT	10	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>Uses a fully reconstructed  $B$  meson as a tag on the recoil side. $\Gamma(D^{*-}\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.67±0.08±0.10</b>	ABDALLAH 04D	DLPH	$e^+e^- \rightarrow Z^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.6 ± 0.3 ± 0.1	<sup>1</sup> BARISH	95	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup>BARISH 95 use  $B(D^0 \rightarrow K^-\pi^+) = (3.91 \pm 0.08 \pm 0.17)\%$  and  $B(D^{*+} \rightarrow D^0\pi^+) = (68.1 \pm 1.0 \pm 1.3)\%$ . $\Gamma(D^{*0}\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_8/\Gamma$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
0.6 ± 0.6 ± 0.1	<sup>1</sup> BARISH	95	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>BARISH 95 use  $B(D^0 \rightarrow K^-\pi^+) = (3.91 \pm 0.08 \pm 0.17)\%$ ,  $B(D^{*+} \rightarrow D^0\pi^+) = (68.1 \pm 1.0 \pm 1.3)\%$ ,  $B(D^{*0} \rightarrow D^0\pi^0) = (63.6 \pm 2.3 \pm 3.3)\%$ . $\Gamma(D^{**}\ell^+\nu_\ell)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma$  $D^{**}$  stands for the sum of the  $D(1^1P_1)$ ,  $D(1^3P_0)$ ,  $D(1^3P_1)$ ,  $D(1^3P_2)$ ,  $D(2^1S_0)$ , and  $D(2^1S_1)$  resonances.  $\ell = e \text{ or } \mu$ , not sum over  $e$  and  $\mu$  modes.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.027±0.005±0.005</b>	63	<sup>1</sup> ALBRECHT	93	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.028	95	<sup>2</sup> BARISH	95	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
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<sup>1</sup>ALBRECHT 93 assumes the GISW model to correct for unseen modes. Using the BHKT model, the result becomes  $0.023 \pm 0.006 \pm 0.004$ . Assumes  $B(D^{*+} \rightarrow D^0\pi^+) = 68.1\%$ ,  $B(D^0 \rightarrow K^-\pi^+) = 3.65\%$ ,  $B(D^0 \rightarrow K^-\pi^+\pi^-\pi^+) = 7.5\%$ . We have taken their average  $e$  and  $\mu$  value.<sup>2</sup>BARISH 95 use  $B(D^0 \rightarrow K^-\pi^+) = (3.91 \pm 0.08 \pm 0.17)\%$ , assume all nonresonant channels are zero, and use GISW model for relative abundances of  $D^{**}$  states. $\Gamma(\bar{D}_1(2420)\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0038±0.0013 OUR AVERAGE</b>	Error includes scale factor of 2.4.		
0.0033±0.0006	<sup>1</sup> ABAZOV	05o	D0 $p\bar{p}$ at 1.96 TeV
0.0074±0.0016	<sup>2</sup> BUSKULIC	97b	ALEP $e^+e^- \rightarrow Z$

• • • We do not use the following data for averages, fits, limits, etc. • • •

seen	<sup>3</sup> BUSKULIC	95b	ALEP	Repl. by BUSKULIC 97b
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<sup>1</sup>Assumes  $B(D_1 \rightarrow D^*\pi) = 1$ ,  $B(D_1 \rightarrow D^*\pi^\pm) = 2/3$ , and  $B(b \rightarrow B) = 0.397$ .<sup>2</sup>BUSKULIC 97b assumes  $B(D_1(2420) \rightarrow D^*\pi) = 1$ ,  $B(D_1(2420) \rightarrow D^*\pi^\pm) = 2/3$ , and  $B(b \rightarrow B) = 0.378 \pm 0.022$ .<sup>3</sup>BUSKULIC 95b reports  $f_B \times B(B \rightarrow \bar{D}_1(2420)^0\ell^+\nu_\ell \text{ anything}) \times B(\bar{D}_1(2420)^0 \rightarrow \bar{D}^*(2010)^-\pi^+) = (2.04 \pm 0.58 \pm 0.34)10^{-3}$ , where  $f_B$  is the production fraction for a single  $B$  charge state. $[\Gamma(D\pi\ell^+\nu_\ell \text{ anything}) + \Gamma(D^*\pi\ell^+\nu_\ell \text{ anything})]/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.026 ± 0.005 OUR AVERAGE</b>	Error includes scale factor of 1.5.		
0.0340±0.0052±0.0032	<sup>1</sup> ABREU	00R	DLPH $e^+e^- \rightarrow Z$
0.0226±0.0029±0.0033	<sup>2</sup> BUSKULIC	97b	ALEP $e^+e^- \rightarrow Z$

<sup>1</sup>Assumes no contribution from  $B_S$  and  $b$  baryons. Further assumes contributions from single pion ( $D\pi$  and  $D^*\pi$ ) states only, allowing isospin conservation to relate the relative  $\pi^0$  and  $\pi^\pm$  rates.<sup>2</sup>BUSKULIC 97b assumes  $B(b \rightarrow B) = 0.378 \pm 0.022$  and uses isospin invariance by assuming that all observed  $D^0\pi^+$ ,  $D^{*0}\pi^+$ ,  $D^+\pi^-$ , and  $D^{*+}\pi^-$  are from  $D^{**}$  states. A correction has been applied to account for the production of  $B_S^0$  and  $\Lambda_b^0$ . $\Gamma(D\pi\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0154±0.0061</b>	ABREU	00R	DLPH $e^+e^- \rightarrow Z$

 $\Gamma(D^*\pi\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0186±0.0038</b>	ABREU	00R	DLPH $e^+e^- \rightarrow Z$

 $\Gamma(\bar{D}_2^*(2460)\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.0044±0.0016</b>		<sup>1</sup> ABAZOV	05o	D0 $p\bar{p}$ at 1.96 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.0065	95	<sup>2</sup> BUSKULIC	97b	ALEP $e^+e^- \rightarrow Z$
not seen		<sup>3</sup> BUSKULIC	95b	ALEP $e^+e^- \rightarrow Z$

<sup>1</sup>Assumes  $B(D_2^* \rightarrow D^*\pi^\pm) = 0.30 \pm 0.06$  and  $B(b \rightarrow B) = 0.397$ .<sup>2</sup>A revised number based on BUSKULIC 97b which assumes  $B(\bar{D}_2^*(2460) \rightarrow D^*\pi^\pm) = 0.20$  and  $B(b \rightarrow B) = 0.378 \pm 0.022$ .<sup>3</sup>BUSKULIC 95b reports  $f_B \times B(B \rightarrow \bar{D}_2^*(2460)^0\ell^+\nu_\ell \text{ anything}) \times B(\bar{D}_2^*(2460)^0 \rightarrow \bar{D}^*(2010)^-\pi^+) \leq 0.81 \times 10^{-3}$  at CL=95%, where  $f_B$  is the production fraction for a single  $B$  charge state. $\Gamma(B \rightarrow \bar{D}_2^*(2460)\ell^+\nu_\ell \text{ anything}) \times B(D_2^*(2460) \rightarrow D^{*-}\pi^+)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.39±0.09±0.12</b>	ABAZOV	05o	D0 $p\bar{p}$ at 1.96 TeV

 $\Gamma(D^{*-}\pi^+\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{16}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>10.0±2.7±2.1</b>	<sup>1</sup> BUSKULIC	95b	ALEP $e^+e^- \rightarrow Z$

<sup>1</sup>BUSKULIC 95b reports  $f_B \times B(B \rightarrow \bar{D}^*(2010)^-\pi^+\ell^+\nu_\ell \text{ anything}) = (3.7 \pm 1.0 \pm 0.7)10^{-3}$ . Above value assumes  $f_B = 0.37 \pm 0.03$ . $\Gamma(\bar{D}\pi^+\pi^-\ell^+\nu_\ell)/\Gamma(\bar{D}\ell^+\nu_\ell)$   $\Gamma_{17}/\Gamma_6$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.7±1.0±0.8</b>	<sup>1</sup> LEES	16	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>Measurement used electrons and muons as leptons. $\Gamma(\bar{D}^*\pi^+\pi^-\ell^+\nu_\ell)/\Gamma(D^*\ell^+\nu_\ell)$   $\Gamma_{18}/\Gamma_9$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.9±0.5±0.4</b>	<sup>1</sup> LEES	16	BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>Measurement used electrons and muons as leptons. $\Gamma(D_s^-\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{19}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7 × 10<sup>-3</sup></b>	90	<sup>1</sup> ALBRECHT	93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>ALBRECHT 93E reports < 0.012 from a measurement of  $[\Gamma(B \rightarrow D_s^-\ell^+\nu_\ell \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ . $\Gamma(D_s^-\ell^+\nu_\ell K^+ \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{20}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5 × 10<sup>-3</sup></b>	90	<sup>1</sup> ALBRECHT	93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>ALBRECHT reports < 0.008 from a measurement of  $[\Gamma(B \rightarrow D_s^-\ell^+\nu_\ell K^+ \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ . $\Gamma(D_s^-\ell^+\nu_\ell K^0 \text{ anything})/\Gamma_{\text{total}}$   $\Gamma_{21}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7 × 10<sup>-3</sup></b>	90	<sup>1</sup> ALBRECHT	93E	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>ALBRECHT reports < 0.012 from a measurement of  $[\Gamma(B \rightarrow D_s^-\ell^+\nu_\ell K^0 \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)]$  assuming  $B(D_s^+ \rightarrow \phi\pi^+) = 0.027$ , which we rescale to our best value  $B(D_s^+ \rightarrow \phi\pi^+) = 4.5 \times 10^{-2}$ .

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

### $\Gamma(X_c \ell^+ \nu_\ell)/\Gamma_{\text{total}}$ $\Gamma_{22}/\Gamma$

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.1065 ± 0.0016 OUR EVALUATION</b>			
<b>0.1058 ± 0.0015 OUR AVERAGE</b>			
0.1064 ± 0.0017 ± 0.0006	<sup>1</sup> AUBERT	10A	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
0.1044 ± 0.0019 ± 0.0022	<sup>2</sup> URQUIJO	07	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1061 ± 0.0016 ± 0.0006	<sup>3</sup> AUBERT	04x	BABR Repl. by AUBERT 10A

- <sup>1</sup> Obtained from a combined fit to the moments of observed spectra in inclusive  $B \rightarrow X_c \ell^+ \nu_\ell$  decay.
- <sup>2</sup> Measured the independent  $B^+$  and  $B^0$  partial branching fractions with electron energy above 0.4 GeV.
- <sup>3</sup> The semileptonic branching ratio,  $|V_{cb}|$  and other heavy-quark parameters are determined from a simultaneous fit to moments of the hadronic-mass and lepton-energy distribution.

### $\Gamma(X_u \ell^+ \nu_\ell)/\Gamma_{\text{total}}$ $\Gamma_{23}/\Gamma$

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.14 ± 0.31 OUR EVALUATION</b>			
2.01 ± 0.15 ± 0.25	<sup>1</sup> LEES	12R	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
2.27 ± 0.26 $^{+0.37}_{-0.33}$	<sup>2</sup> AUBERT	06H	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
2.53 ± 0.24 ± 0.24	<sup>3</sup> AUBERT,B	05x	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
2.80 ± 0.52 ± 0.41	<sup>4</sup> LIMOSA NI	05	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
1.77 ± 0.29 ± 0.38	<sup>5</sup> BORNHEIM	02	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.963 ± 0.173 ± 0.159	<sup>6</sup> URQUIJO	10	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
1.18 ± 0.09 ± 0.07	<sup>7</sup> AUBERT	08As	BABR Repl. by LEES 12R
2.24 ± 0.27 ± 0.47	<sup>8,9</sup> AUBERT	04i	BABR Repl. by AUBERT,B 05x

- <sup>1</sup> Measures several partial branching fractions in different phase space regions. The most precise result on the full branching fraction is obtained in the region for lepton momentum in  $B$  rest frame  $p_\ell^+ > 1$  GeV/c, where the measured partial branching fraction is  $\Delta B = (1.80 \pm 0.13 \pm 0.15) \times 10^{-3}$ . The acceptance in that region is reported in a private communication by the Authors to be 0.894. The corresponding  $|V_{ub}|$  from the BLNP method is  $(4.28 \pm 0.15 \pm 0.18 \pm 0.19) \times 10^{-3}$ , where the last uncertainty comes from theoretical prediction.
- <sup>2</sup> Obtained from the partial rate  $\Delta B = (0.572 \pm 0.041 \pm 0.065) \times 10^{-3}$  for the electron momentum interval of 2.0–2.6 GeV/c based on BLNP method.
- <sup>3</sup> Determined from the partial rate  $\Delta B = (4.41 \pm 0.42 \pm 0.42) \times 10^{-4}$  measured for electron energy  $> 2$  GeV and hadronic mass squared  $< 3.5$  GeV<sup>2</sup>, and calculated acceptance 0.174 in that region. The  $V_{ub}$  is measured as  $(4.41 \pm 0.30^{+0.65}_{-0.47} \pm 0.28) \times 10^{-3}$ .
- <sup>4</sup> Uses electrons in the momentum interval 1.9–2.6 GeV/c in the center-of-mass frame. The  $V_{ub}$  is found to be  $(5.08 \pm 0.47^{+0.49}_{-0.48}) \times 10^{-3}$ .
- <sup>5</sup> BORNHEIM 02 uses the observed yield of leptons from semileptonic  $B$  decays in the end-point momentum interval 2.2–2.6 GeV/c with recent CLEO-2 data on  $B \rightarrow X_S \gamma$ . The  $V_{ub}$  is found to be  $(4.08 \pm 0.34 \pm 0.53) \times 10^{-3}$ .
- <sup>6</sup> Uses a multivariate analysis method and requires lepton momentum in the  $B$  rest frame,  $p_\ell^{*B} > 1.0$  GeV/c.
- <sup>7</sup> Measures several partial branching fractions in different phase space regions. The most precise result is obtained in the region for hadronic mass  $M_X < 1.55$  GeV/c<sup>2</sup>, and is  $\Delta B = (1.18 \pm 0.09 \pm 0.07) \times 10^{-3}$ . The corresponding  $|V_{ub}|$  from the BLNP method is  $(4.27 \pm 0.16 \pm 0.13 \pm 0.30) \times 10^{-3}$ , where the last uncertainty comes from the theoretical prediction of the partial rate in the given phase-space region.
- <sup>8</sup> Used BaBar measurement of Semileptonic branching fraction  $B(B \rightarrow X \ell \nu_\ell) = (10.87 \pm 0.18 \pm 0.30)\%$  to convert the ratio of rates to branching fraction.
- <sup>9</sup> The third error includes the systematics and theoretical errors summed in quadrature.

### $\Gamma(X_u \ell^+ \nu_\ell)/\Gamma(\ell^+ \nu_\ell \text{ anything})$ $\Gamma_{23}/\Gamma_3$

$\ell$  denotes  $e$  or  $\mu$ , not the sum. These experiments measure this ratio in very limited momentum intervals.

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.06 ± 0.25 ± 0.42</b>			<sup>1</sup> AUBERT	04i	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
			<sup>2</sup> ALBRECHT	94c	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
			<sup>3</sup> BARTELT	93B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
		107	<sup>4</sup> ALBRECHT	91c	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
		77	<sup>5</sup> ALBRECHT	90	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
		41	<sup>6</sup> FULTON	90	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
		76	<sup>7</sup> BEHREND	87	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
<4.0	90		CHEN	84	CLEO Direct $e$ at $\Upsilon(4S)$
<4.0	90		KLOPFEN...	83B	CUSB Direct $e$ at $\Upsilon(4S)$
<5.5	90				
<sup>1</sup> The third error includes the systematics and theoretical errors summed in quadrature.					

- <sup>2</sup> ALBRECHT 94c find  $\Gamma(b \rightarrow c)/\Gamma(b \rightarrow \text{all}) = 0.99 \pm 0.02 \pm 0.04$ .
- <sup>3</sup> BARTELT 93B (CLEO II) measures an excess of  $107 \pm 15 \pm 11$  leptons in the lepton momentum interval 2.3–2.6 GeV/c which is attributed to  $b \rightarrow u \ell \nu_\ell$ . This corresponds to a model-dependent partial branching ratio  $\Delta B_{ub}$  between  $(1.15 \pm 0.16 \pm 0.15) \times 10^{-4}$ , as evaluated using the KS model (KOERNER 88), and  $(1.54 \pm 0.22 \pm 0.20) \times 10^{-4}$  using the ACCMM model (ARTUSO 93). The corresponding values of  $|V_{ub}|/|V_{cb}|$  are  $0.056 \pm 0.006$  and  $0.076 \pm 0.008$ , respectively.
- <sup>4</sup> ALBRECHT 91c result supersedes ALBRECHT 90. Two events are fully reconstructed providing evidence for the  $b \rightarrow u$  transition. Using the model of ALTARELLI 82, they obtain  $|V_{ub}/V_{cb}| = 0.11 \pm 0.012$  from 77 leptons in the 2.3–2.6 GeV momentum range.
- <sup>5</sup> ALBRECHT 90 observes 41 ± 10 excess  $e$  and  $\mu$  (lepton) events in the momentum interval  $p = 2.3\text{--}2.6$  GeV signaling the presence of the  $b \rightarrow u$  transition. The events correspond to a model-dependent measurement of  $|V_{ub}/V_{cb}| = 0.10 \pm 0.01$ .
- <sup>6</sup> FULTON 90 observe 76 ± 20 excess  $e$  and  $\mu$  (lepton) events in the momentum interval  $p = 2.4\text{--}2.6$  GeV signaling the presence of the  $b \rightarrow u$  transition. The average branching ratio,  $(1.8 \pm 0.4 \pm 0.3) \times 10^{-4}$ , corresponds to a model-dependent measurement of approximately  $|V_{ub}/V_{cb}| = 0.1$  using  $B(b \rightarrow c \ell \nu) = 10.2 \pm 0.2 \pm 0.7\%$ .
- <sup>7</sup> The quoted possible limits range from 0.018 to 0.04 for the ratio, depending on which model or momentum range is chosen. We select the most conservative limit they have calculated. This corresponds to a limit on  $|V_{ub}|/|V_{cb}| < 0.20$ . While the endpoint technique employed is more robust than their previous results in CHEN 84, these results do not provide a numerical improvement in the limit.

### $\Gamma(K^+ \ell^+ \nu_\ell \text{ anything})/\Gamma(\ell^+ \nu_\ell \text{ anything})$ $\Gamma_{24}/\Gamma_3$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.58 ± 0.05 OUR AVERAGE</b>			
0.594 ± 0.021 ± 0.056	ALBRECHT	94c	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
0.54 ± 0.07 ± 0.06	<sup>1</sup> ALAM	87B	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALAM 87B measurement relies on lepton-kaon correlations.			

### $\Gamma(K^- \ell^+ \nu_\ell \text{ anything})/\Gamma(\ell^+ \nu_\ell \text{ anything})$ $\Gamma_{25}/\Gamma_3$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.092 ± 0.035 OUR AVERAGE</b>			
0.086 ± 0.011 ± 0.044	ALBRECHT	94c	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
0.10 ± 0.05 ± 0.02	<sup>1</sup> ALAM	87B	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALAM 87B measurement relies on lepton-kaon correlations.			

### $\Gamma(K^0/\bar{K}^0 \ell^+ \nu_\ell \text{ anything})/\Gamma(\ell^+ \nu_\ell \text{ anything})$ $\Gamma_{26}/\Gamma_3$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.42 ± 0.05 OUR AVERAGE</b>			
0.452 ± 0.038 ± 0.056	<sup>1</sup> ALBRECHT	94c	ARG $e^+ e^- \rightarrow \Upsilon(4S)$
0.39 ± 0.06 ± 0.04	<sup>2</sup> ALAM	87B	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 94c assume a $K^0/\bar{K}^0$ multiplicity twice that of $K_S^0$ .			
<sup>2</sup> ALAM 87B measurement relies on lepton-kaon correlations.			

### $\Gamma(\bar{D}^+ \tau^+ \nu_\tau)/\Gamma(\bar{D} \ell^+ \nu_\ell)$ $\Gamma_{27}/\Gamma_6$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>41 ± 5 OUR AVERAGE</b>			
37.5 ± 6.4 ± 2.6	<sup>1,2</sup> HUSCHLE	15	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
44.0 ± 5.8 ± 4.2	<sup>1,2</sup> LEES	12D	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.16 ± 11.7 ± 5.2	<sup>1</sup> AUBERT	08N	BABR Repl. by LEES 12D
<sup>1</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side.			
<sup>2</sup> Uses $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$ and $\tau^+ \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$ and $e^+$ or $\mu^+$ as $\ell^+$ . Obtained from simultaneous fit to $B^+$ and $B^0$ assuming isospin symmetry.			

### $\Gamma(D^+ \tau^+ \nu_\tau)/\Gamma(D^+ \ell^+ \nu_\ell)$ $\Gamma_{28}/\Gamma_9$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>31.8 ± 2.4 OUR AVERAGE</b>			
29.3 ± 3.8 ± 1.5	<sup>1</sup> HUSCHLE	15	BELL $e^+ e^- \rightarrow \Upsilon(4S)$
33.2 ± 2.4 ± 1.8	<sup>1</sup> LEES	12D	BABR $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
29.7 ± 5.6 ± 1.8	<sup>2</sup> AUBERT	08N	BABR Repl. by LEES 12D
<sup>1</sup> Uses $\tau^+ \rightarrow e^+ \nu_e \bar{\nu}_\tau$ and $\tau^+ \rightarrow \mu^+ \nu_\mu \bar{\nu}_\tau$ and $e^+$ or $\mu^+$ as $\ell^+$ . Obtained from simultaneous fit to $B^+$ and $B^0$ assuming isospin symmetry. Uses a fully reconstructed $B$ meson as a tag on the recoil side.			
<sup>2</sup> Uses a fully reconstructed $B$ meson as a tag on the recoil side. The results are normalized to the $B^+$ decay rate.			

### $\langle n_c \rangle$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.10 ± 0.05</b>	<sup>1</sup> GIBBONS	97B	CLE2 $e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.98 ± 0.16 ± 0.12	<sup>2</sup> ALAM	87B	CLEO $e^+ e^- \rightarrow \Upsilon(4S)$



See key on page 601

## Meson Particle Listings

 $B^\pm/B^0$  ADMIXTURE

<sup>1</sup> GIBBONS 97B from charm counting using  $B(D_s^+ \rightarrow \phi\pi) = 0.036 \pm 0.009$  and  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.044 \pm 0.006$ .

<sup>2</sup> From the difference between  $K^-$  and  $K^+$  widths. ALAM 87B measurement relies on lepton-kaon correlations. It does not consider the possibility of  $B\bar{B}$  mixing. We have thus removed it from the average.

$\Gamma(D^\pm \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{29}/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.229<math>\pm</math>0.013 OUR AVERAGE</b>				
0.228 $\pm$ 0.012 $\pm$ 0.006		<sup>1</sup> GIBBONS 97B	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
0.24 $\pm$ 0.04 $\pm$ 0.01		<sup>2</sup> BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
0.22 $\pm$ 0.05 $\pm$ 0.01		<sup>3</sup> ALBRECHT 91H	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.20 $\pm$ 0.05 $\pm$ 0.01	20k	<sup>4</sup> BORTOLETTO87	CLEO	Sup. by BORTOLETTO 92

<sup>1</sup> GIBBONS 97B reports  $[\Gamma(B \rightarrow D^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^- 2\pi^+)] = 0.0216 \pm 0.0008 \pm 0.00082$  which we divide by our best value  $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> BORTOLETTO 92 reports  $[\Gamma(B \rightarrow D^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^- 2\pi^+)] = 0.0226 \pm 0.0030 \pm 0.0018$  which we divide by our best value  $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> ALBRECHT 91H reports  $[\Gamma(B \rightarrow D^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^- 2\pi^+)] = 0.0209 \pm 0.0027 \pm 0.0040$  which we divide by our best value  $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> BORTOLETTO 87 reports  $[\Gamma(B \rightarrow D^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^- 2\pi^+)] = 0.019 \pm 0.004 \pm 0.002$  which we divide by our best value  $B(D^+ \rightarrow K^- 2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D^0/\bar{D}^0 \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{30}/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.618<math>\pm</math>0.029 OUR AVERAGE</b>				
0.639 $\pm$ 0.024 $\pm$ 0.006 -0.007		<sup>1</sup> GIBBONS 97B	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
0.59 $\pm$ 0.05 $\pm$ 0.01		<sup>2</sup> BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
0.494 $\pm$ 0.074 $\pm$ 0.005		<sup>3</sup> ALBRECHT 91H	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.53 $\pm$ 0.07 $\pm$ 0.01	21k	<sup>4</sup> BORTOLETTO87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
0.61 $\pm$ 0.18 $\pm$ 0.01		<sup>5</sup> GREEN 83	CLEO	Repl. by BORTOLETTO 87

<sup>1</sup> GIBBONS 97B reports  $[\Gamma(B \rightarrow D^0/\bar{D}^0 \text{ anything})/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^- \pi^+)] = 0.0251 \pm 0.0006 \pm 0.00075$  which we divide by our best value  $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

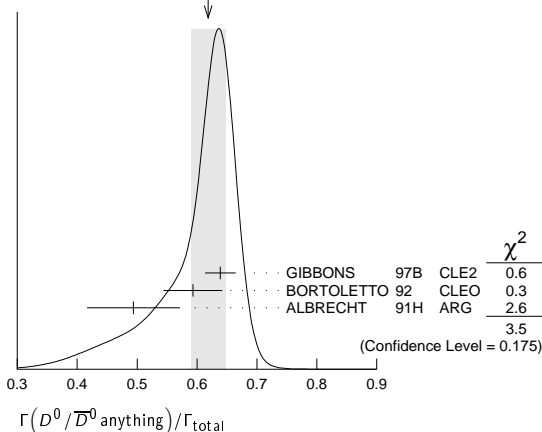
<sup>2</sup> BORTOLETTO 92 reports  $[\Gamma(B \rightarrow D^0/\bar{D}^0 \text{ anything})/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^- \pi^+)] = 0.0233 \pm 0.0012 \pm 0.0014$  which we divide by our best value  $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> ALBRECHT 91H reports  $[\Gamma(B \rightarrow D^0/\bar{D}^0 \text{ anything})/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^- \pi^+)] = 0.0194 \pm 0.0015 \pm 0.0025$  which we divide by our best value  $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> BORTOLETTO 87 reports  $[\Gamma(B \rightarrow D^0/\bar{D}^0 \text{ anything})/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^- \pi^+)] = 0.0210 \pm 0.0015 \pm 0.0021$  which we divide by our best value  $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>5</sup> GREEN 83 reports  $[\Gamma(B \rightarrow D^0/\bar{D}^0 \text{ anything})/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^- \pi^+)] = 0.024 \pm 0.006 \pm 0.004$  which we divide by our best value  $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

WEIGHTED AVERAGE  
0.618 $\pm$ 0.029 (Error scaled by 1.3)



$\Gamma(D^*(2010)^\pm \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{31}/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.225<math>\pm</math>0.015 OUR AVERAGE</b>				
0.247 $\pm$ 0.019 $\pm$ 0.01		<sup>1</sup> GIBBONS 97B	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
0.205 $\pm$ 0.019 $\pm$ 0.007		<sup>2</sup> ALBRECHT 96D	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
0.230 $\pm$ 0.028 $\pm$ 0.009		<sup>3</sup> BORTOLETTO92	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.283 $\pm$ 0.053 $\pm$ 0.002		<sup>4</sup> ALBRECHT 91H	ARG	Sup. by ALBRECHT 96D
0.22 $\pm$ 0.04 $\pm$ 0.07 -0.04	5200	<sup>5</sup> BORTOLETTO87	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
0.27 $\pm$ 0.06 $\pm$ 0.08 -0.06	510	<sup>6</sup> CSORNA 85	CLEO	Repl. by BORTOLETTO 87

<sup>1</sup> GIBBONS 97B reports  $B(B \rightarrow D^*(2010)^\pm \text{ anything}) = 0.239 \pm 0.015 \pm 0.014 \pm 0.009$  using CLEO measured  $D$  and  $D^*$  branching fractions. We rescale to our PDG 96 values of  $D$  and  $D^*$  branching ratios. Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ALBRECHT 96D reports  $B(B \rightarrow D^*(2010)^\pm \text{ anything}) = 0.196 \pm 0.019$  using CLEO measured  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.681 \pm 0.01 \pm 0.013$ ,  $B(D^0 \rightarrow K^- \pi^+) = 0.0401 \pm 0.0014$ ,  $B(D^0 \rightarrow K^- \pi^+ \pi^+ \pi^-) = 0.081 \pm 0.005$ . We rescale to our PDG 96 values of  $D$  and  $D^*$  branching ratios. Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> BORTOLETTO 92 reports  $B(B \rightarrow D^*(2010)^\pm \text{ anything}) = 0.25 \pm 0.03 \pm 0.04$  using MARK II  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.57 \pm 0.06$  and  $B(D^0 \rightarrow K^- \pi^+) = 0.042 \pm 0.008$ . We rescale to our PDG 96 values of  $D$  and  $D^*$  branching ratios. Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> ALBRECHT 91H reports  $0.348 \pm 0.060 \pm 0.035$  from a measurement of  $[\Gamma(B \rightarrow D^*(2010)^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D^*(2010)^+ \rightarrow D^0 \pi^+)]$  assuming  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.55 \pm 0.04$ , which we rescale to our best value  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (67.7 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Uses the PDG 90  $B(D^0 \rightarrow K^- \pi^+) = 0.0371 \pm 0.0025$ .

<sup>5</sup> BORTOLETTO 87 uses old MARK III (BALTRUSAITIS 86E) branching ratios  $B(D^0 \rightarrow K^- \pi^+) = 0.056 \pm 0.004 \pm 0.003$  and also assumes  $B(D^*(2010)^+ \rightarrow D^0 \pi^+) = 0.60 \pm 0.08 \pm 0.15$ . The product branching ratio for  $B(B \rightarrow D^*(2010)^+) B(D^*(2010)^+ \rightarrow D^0 \pi^+)$  is  $0.13 \pm 0.02 \pm 0.012$ . Superseded by BORTOLETTO 92.

<sup>6</sup> V-A momentum spectrum used to extrapolate below  $p = 1$  GeV. We correct the value assuming  $B(D^0 \rightarrow K^- \pi^+) = 0.042 \pm 0.006$  and  $B(D^+ \rightarrow D^0 \pi^+) = 0.6 \pm 0.08 \pm 0.15$ . The product branching fraction is  $B(B \rightarrow D^*(2010)^+) B(D^*(2010)^+ \rightarrow D^0 \pi^+) = (68 \pm 15 \pm 9) \times 10^{-4}$ .

$\Gamma(D^*(2007)^0 \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{32}/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.260<math>\pm</math>0.023<math>\pm</math>0.015</b>				
		<sup>1</sup> GIBBONS 97B	CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> GIBBONS 97B reports  $B(B \rightarrow D^*(2007)^0 \text{ anything}) = 0.247 \pm 0.012 \pm 0.018 \pm 0.018$  using CLEO measured  $D$  and  $D^*$  branching fractions. We rescale to our PDG 96 values of  $D$  and  $D^*$  branching ratios. Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D_s^\pm \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{33}/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.083<math>\pm</math>0.008 OUR AVERAGE</b>				
0.089 $\pm$ 0.010 $\pm$ 0.008		<sup>1</sup> ARTUSO 05B	CLE2	$e^+e^- \rightarrow \Upsilon(5S)$
0.087 $\pm$ 0.005 $\pm$ 0.008		<sup>2</sup> AUBERT 02G	BABR	$e^+e^- \rightarrow \Upsilon(4S)$
0.065 $\pm$ 0.011 $\pm$ 0.006		<sup>3</sup> ALBRECHT 92G	ARG	$e^+e^- \rightarrow \Upsilon(4S)$
0.068 $\pm$ 0.010 $\pm$ 0.006	257	<sup>4</sup> BORTOLETTO90	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
0.085 $\pm$ 0.022 $\pm$ 0.008		<sup>5</sup> HAAS 86	CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.094 $\pm$ 0.007 $\pm$ 0.008		<sup>6</sup> GIBAUT 96	CLE2	Repl. by ARTUSO 05B
0.094 $\pm$ 0.024 $\pm$ 0.008		<sup>7</sup> ALBRECHT 87H	ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ARTUSO 05B reports  $0.0905 \pm 0.0025 \pm 0.0140$  from a measurement of  $[\Gamma(B \rightarrow D_s^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^\pm \rightarrow \phi \pi^\pm)]$  assuming  $B(D_s^\pm \rightarrow \phi \pi^\pm) = (4.4 \pm 0.5) \times 10^{-2}$ , which we rescale to our best value  $B(D_s^\pm \rightarrow \phi \pi^\pm) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> AUBERT 02G reports  $[\Gamma(B \rightarrow D_s^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^\pm \rightarrow \phi \pi^\pm)] = 0.00393 \pm 0.00007 \pm 0.00021$  which we divide by our best value  $B(D_s^\pm \rightarrow \phi \pi^\pm) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> ALBRECHT 92G reports  $[\Gamma(B \rightarrow D_s^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^\pm \rightarrow \phi \pi^\pm)] = 0.00292 \pm 0.00039 \pm 0.00031$  which we divide by our best value  $B(D_s^\pm \rightarrow \phi \pi^\pm) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> BORTOLETTO 90 reports  $[\Gamma(B \rightarrow D_s^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^\pm \rightarrow \phi \pi^\pm)] = 0.00306 \pm 0.00047$  which we divide by our best value  $B(D_s^\pm \rightarrow \phi \pi^\pm) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>5</sup> HAAS 86 reports  $[\Gamma(B \rightarrow D_s^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^\pm \rightarrow \phi \pi^\pm)] = 0.0038 \pm 0.0010$  which we divide by our best value  $B(D_s^\pm \rightarrow \phi \pi^\pm) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. 64  $\pm$  22% decays are 2-body.

<sup>6</sup> GIBAUT 96 reports  $0.1211 \pm 0.0039 \pm 0.0088$  from a measurement of  $[\Gamma(B \rightarrow D_s^\pm \text{ anything})/\Gamma_{\text{total}}] \times [B(D_s^\pm \rightarrow \phi \pi^\pm)]$  assuming  $B(D_s^\pm \rightarrow \phi \pi^\pm) = 0.035$ , which



# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

### $\Gamma(\psi(2S) \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{50}/\Gamma$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.00307 ± 0.00021 OUR AVERAGE</b>				
0.00297 ± 0.00020 ± 0.00020		AUBERT	03F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
0.00316 ± 0.00014 ± 0.00028		<sup>1</sup> ANDERSON	02 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
0.0046 ± 0.0017 ± 0.0011	8	ALBRECHT	87D ARG	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0034 ± 0.0004 ± 0.0003	240	<sup>2</sup> BALEST	95B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Also reports the measurement of  $\psi(2S) \rightarrow \ell^+ \ell^-$  polarization produced directly from B decay.

<sup>2</sup> BALEST 95B assume PDG 1994 values for sub mode branching ratios. They find  $B(B \rightarrow \psi(2S)X, \psi(2S) \rightarrow \ell^+ \ell^-) = 0.30 \pm 0.05 \pm 0.04$  and  $B(B \rightarrow \psi(2S)X, \psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = 0.37 \pm 0.05 \pm 0.05$ . Weighted average is quoted for  $B(B \rightarrow \psi(2S)X)$ .

### $\Gamma(\chi_{c1}(1P) \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{51}/\Gamma$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.00386 ± 0.00027 OUR AVERAGE</b>				
0.00367 ± 0.00035 ± 0.00044		AUBERT	03F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
0.00363 ± 0.00022 ± 0.00034		<sup>1</sup> ABE	02L BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.00435 ± 0.00029 ± 0.00040		ANDERSON	02 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0033 ± 0.0004 ± 0.0001		<sup>2</sup> CHEN	01 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
0.0040 ± 0.0006 ± 0.0004	112	<sup>3</sup> BALEST	95B CLE2	Repl. by CHEN 01
0.0105 ± 0.0035 ± 0.0025		<sup>4</sup> ALBRECHT	92E ARG	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ABE 02L uses PDG 01 values for  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-)$  and  $B(\chi_{c1,c2} \rightarrow J/\psi(1S)\gamma)$ .

<sup>2</sup> CHEN 01 reports  $0.00414 \pm 0.00031 \pm 0.00040$  from a measurement of  $[\Gamma(B \rightarrow \chi_{c1}(1P) \text{ anything})/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> BALEST 95B assume  $B(\chi_{c1}(1P) \rightarrow J/\psi(1S)\gamma) = (27.3 \pm 1.6) \times 10^{-2}$ , the PDG 1994 value. Fit to  $\psi$ -photon invariant mass distribution allows for a  $\chi_{c1}(1P)$  and a  $\chi_{c2}(1P)$  component.

<sup>4</sup> ALBRECHT 92E assumes no  $\chi_{c2}(1P)$  production.

### $\Gamma(\chi_{c1}(1P) \text{ (direct) anything})/\Gamma_{\text{total}}$ $\Gamma_{52}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.00324 ± 0.00025 OUR AVERAGE</b>			
0.00341 ± 0.00035 ± 0.00042	AUBERT	03F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
0.00332 ± 0.00022 ± 0.00034	<sup>1</sup> ABE	02L BELL	$e^+e^- \rightarrow \Upsilon(4S)$
0.0031 ± 0.0004 ± 0.0001	<sup>2</sup> CHEN	01 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.0037 ± 0.0007	<sup>3</sup> BALEST	95B CLE2	Repl. by CHEN 01

<sup>1</sup> ABE 02L uses PDG 01 values for  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-)$  and  $B(\chi_{c1,c2} \rightarrow J/\psi(1S)\gamma)$ .

<sup>2</sup> CHEN 01 reports  $0.00383 \pm 0.00031 \pm 0.00040$  from a measurement of  $[\Gamma(B \rightarrow \chi_{c1}(1P) \text{ (direct) anything})/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = 0.273 \pm 0.016$ , which we rescale to our best value  $B(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)) = (33.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> BALEST 95B assume PDG 1994 values.  $J/\psi(1S)$  mesons are reconstructed in the  $e^+e^-$  and  $\mu^+\mu^-$  modes. The  $B \rightarrow \chi_{c1}(1P)X$  branching ratio contains  $\chi_{c1}(1P)$  mesons directly from B decays and also from feeddown through  $\psi(2S) \rightarrow \chi_{c1}(1P)\gamma$ . Using the measured inclusive rates, BALEST 95B corrects for the feeddown and finds the  $B \rightarrow \chi_{c1}(1P) \text{ (direct) } X$  branching ratio.

### $\Gamma(\chi_{c2}(1P) \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{53}/\Gamma$

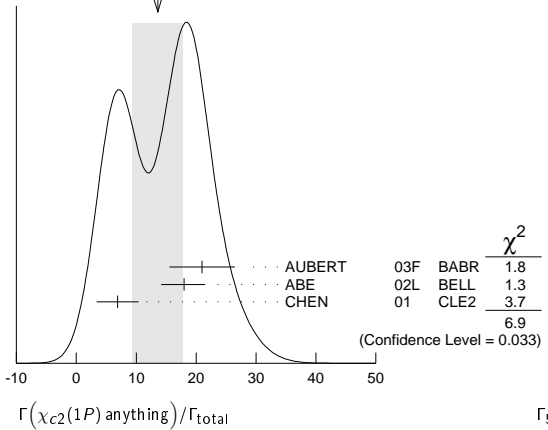
VALUE (units $10^{-4}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>14 ± 4 OUR AVERAGE</b>					Error includes scale factor of 1.9. See the ideogram below.
21.0 ± 4.5 ± 3.1			AUBERT	03F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
18.0 <sup>+2.3</sup> <sub>-2.8</sub> ± 2.6			<sup>1</sup> ABE	02L BELL	$e^+e^- \rightarrow \Upsilon(4S)$
6.9 ± 3.5 ± 0.3			<sup>2</sup> CHEN	01 CLE2	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<38	90	35	<sup>3</sup> BALEST	95B CLE2	Repl. by CHEN 01

<sup>1</sup> ABE 02L uses PDG 01 values for  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-)$  and  $B(\chi_{c1,c2} \rightarrow J/\psi(1S)\gamma)$ .

<sup>2</sup> CHEN 01 reports  $(9.8 \pm 4.8 \pm 1.5) \times 10^{-4}$  from a measurement of  $[\Gamma(B \rightarrow \chi_{c2}(1P) \text{ anything})/\Gamma_{\text{total}}] \times [B(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S))]$  assuming  $B(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)) = 0.135 \pm 0.011$ , which we rescale to our best value  $B(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)) = (19.2 \pm 0.7) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>3</sup> BALEST 95B assume  $B(\chi_{c2}(1P) \rightarrow J/\psi(1S)\gamma) = (13.5 \pm 1.1) \times 10^{-2}$ , the PDG 1994 value.  $J/\psi(1S)$  mesons are reconstructed in the  $e^+e^-$  and  $\mu^+\mu^-$  modes, and PDG 1994 branching fractions are used. If interpreted as signal, the  $35 \pm 13$  events correspond to  $B(B \rightarrow \chi_{c2}(1P)X) = (0.25 \pm 0.10 \pm 0.03) \times 10^{-2}$ .

WEIGHTED AVERAGE  
14 ± 4 (Error scaled by 1.9)



### $\Gamma(\chi_{c2}(1P) \text{ (direct) anything})/\Gamma_{\text{total}}$ $\Gamma_{54}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.00165 ± 0.00031 OUR AVERAGE</b>			
0.00190 ± 0.00045 ± 0.00029	AUBERT	03F BABR	$e^+e^- \rightarrow \Upsilon(4S)$
0.00153 ± 0.00023 ± 0.00027	<sup>1</sup> ABE	02L BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ABE 02L uses PDG 01 values for  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-)$  and  $B(\chi_{c1,c2} \rightarrow J/\psi(1S)\gamma)$ .

### $\Gamma(\eta_c(1S) \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{55}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.009</b>	90	<sup>1</sup> BALEST	95B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> BALEST 95B assume PDG 1994 values for sub mode branching ratios.  $J/\psi(1S)$  mesons are reconstructed in  $J/\psi(1S) \rightarrow e^+e^-$  and  $J/\psi(1S) \rightarrow \mu^+\mu^-$ . Search region  $2960 < m_{\eta_c(1S)} < 3010$  MeV/ $c^2$ .

### $\Gamma(KX(3872), X \rightarrow D^0 \bar{D}^0 \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{56}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.22 ± 0.31 ± 0.23</b>	<sup>1</sup> GOKHROO	06 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Measure the near-threshold enhancements in the  $(D^0 \bar{D}^0 \pi^0)$  system at a mass  $3875.2 \pm 0.7^{+0.3}_{-1.6} \pm 0.8$  MeV/ $c^2$ .

### $\Gamma(KX(3872), X \rightarrow D^{*0} D^0)/\Gamma_{\text{total}}$ $\Gamma_{57}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.80 ± 0.20 ± 0.10</b>	AUSHEV	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

### $\Gamma(KX(3940), X \rightarrow D^{*0} D^0)/\Gamma_{\text{total}}$ $\Gamma_{58}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.67</b>	90	AUSHEV	10 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

### $\Gamma(KX(3915), X \rightarrow \omega J/\psi)/\Gamma_{\text{total}}$ $\Gamma_{59}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>7.1 ± 1.3 ± 3.1</b>	<sup>1</sup> CHOI	05 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> CHOI 05 reports the observation of a near-threshold enhancement in the  $\omega J/\psi$  mass spectrum in exclusive  $B \rightarrow K\omega J/\psi$ . The new state, denoted as  $X(3915)$ , is measured to have a mass of  $3943 \pm 11 \pm 13$  GeV/ $c^2$  and a width  $\Gamma = 87 \pm 22 \pm 26$  MeV.

### $\Gamma(K^\pm \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{60}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.789 ± 0.025 OUR AVERAGE</b>			
0.82 ± 0.01 ± 0.05	ALBRECHT	94C ARG	$e^+e^- \rightarrow \Upsilon(4S)$
0.775 ± 0.015 ± 0.025	<sup>1</sup> ALBRECHT	93I ARG	$e^+e^- \rightarrow \Upsilon(4S)$
0.85 ± 0.07 ± 0.09	ALAM	87B CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
seen	<sup>2</sup> BRODY	82 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$
seen	<sup>3</sup> GIANNINI	82 CUSB	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 93I value is not independent of the sum of  $B \rightarrow K^+ \text{ anything}$  and  $B \rightarrow K^- \text{ anything}$  ALBRECHT 94C values.

<sup>2</sup> Assuming  $\Upsilon(4S) \rightarrow B\bar{B}$ , a total of  $3.38 \pm 0.34 \pm 0.68$  kaons per  $\Upsilon(4S)$  decay is found (the second error is systematic). In the context of the standard B-decay model, this leads to a value for  $(b\text{-quark} \rightarrow c\text{-quark})/(b\text{-quark} \rightarrow \text{all})$  of  $1.09 \pm 0.33 \pm 0.13$ .

<sup>3</sup> GIANNINI 82 at CESR-CUSB observed  $1.58 \pm 0.35$   $K^0$  per hadronic event much higher than  $0.82 \pm 0.10$  below threshold. Consistent with predominant  $b \rightarrow cX$  decay.

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

### $\Gamma(K^+ \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{61}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.66 ± 0.05</b>	<sup>1</sup> ALBRECHT	94c	ARG $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.620 ± 0.013 ± 0.038	<sup>2</sup> ALBRECHT	94c	ARG $e^+e^- \rightarrow \Upsilon(4S)$
0.66 ± 0.05 ± 0.07	<sup>2</sup> ALAM	87b	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Measurement relies on lepton-kaon correlations. It is for the weak decay vertex and does not include mixing of the neutral  $B$  meson. Mixing effects were corrected for by assuming a mixing parameter  $r$  of  $(18.1 \pm 4.3)\%$ .

<sup>2</sup> Measurement relies on lepton-kaon correlations. It includes production through mixing of the neutral  $B$  meson.

### $\Gamma(K^- \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{62}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.13 ± 0.04</b>	<sup>1</sup> ALBRECHT	94c	ARG $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.165 ± 0.011 ± 0.036	<sup>2</sup> ALBRECHT	94c	ARG $e^+e^- \rightarrow \Upsilon(4S)$
0.19 ± 0.05 ± 0.02	<sup>2</sup> ALAM	87b	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Measurement relies on lepton-kaon correlations. It is for the weak decay vertex and does not include mixing of the neutral  $B$  meson. Mixing effects were corrected for by assuming a mixing parameter  $r$  of  $(18.1 \pm 4.3)\%$ .

<sup>2</sup> Measurement relies on lepton-kaon correlations. It includes production through mixing of the neutral  $B$  meson.

### $\Gamma(K^0/\bar{K}^0 \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{63}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.64 ± 0.04 OUR AVERAGE</b>			
0.642 ± 0.010 ± 0.042	<sup>1</sup> ALBRECHT	94c	ARG $e^+e^- \rightarrow \Upsilon(4S)$
0.63 ± 0.06 ± 0.06	ALAM	87b	CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 94c assume a  $K^0/\bar{K}^0$  multiplicity twice that of  $K_S^0$ .

### $\Gamma(K^*(892)^\pm \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{64}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.182 ± 0.054 ± 0.024</b>	ALBRECHT	94j	ARG $e^+e^- \rightarrow \Upsilon(4S)$

### $\Gamma(K^*(892)^0/\bar{K}^*(892)^0 \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{65}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.146 ± 0.016 ± 0.020</b>	ALBRECHT	94j	ARG $e^+e^- \rightarrow \Upsilon(4S)$

### $\Gamma(K^*(892)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{66}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>4.24 ± 0.54 ± 0.32</b>		<sup>1</sup> COAN	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<150	90	<sup>2</sup> LESIAK	92	CBAL $e^+e^- \rightarrow \Upsilon(4S)$
< 24	90	ALBRECHT	88h	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> An average of  $B(B^+ \rightarrow K^*(892)^+\gamma)$  and  $B(B^0 \rightarrow K^*(892)^0\gamma)$  measurements reported in COAN 00 by assuming full correlated systematic errors.

<sup>2</sup> LESIAK 92 set a limit on the inclusive process  $B(b \rightarrow s\gamma) < 2.8 \times 10^{-3}$  at 90% CL for the range of masses of 892–2045 MeV, independent of assumptions about s-quark hadronization.

### $\Gamma(\eta K\gamma)/\Gamma_{\text{total}}$ $\Gamma_{67}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>8.5 ± 1.3 ± 1.2</b> <b>0.9</b>	<sup>1</sup> NISHIDA	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>  $m_{\eta K} < 2.4$  GeV/ $c^2$

### $\Gamma(K_1(1400)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{68}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;12.7 × 10<sup>-5</sup></b>	90	<sup>1</sup> COAN	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1.6 × 10 <sup>-3</sup>	90	<sup>2</sup> LESIAK	92	CBAL $e^+e^- \rightarrow \Upsilon(4S)$
< 4.1 × 10 <sup>-4</sup>	90	ALBRECHT	88h	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

<sup>2</sup> LESIAK 92 set a limit on the inclusive process  $B(b \rightarrow s\gamma) < 2.8 \times 10^{-3}$  at 90% CL for the range of masses of 892–2045 MeV, independent of assumptions about s-quark hadronization.

### $\Gamma(K_2^*(1430)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{69}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.66 ± 0.59 ± 0.13</b> <b>0.53</b>		<sup>1</sup> COAN	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 83	90	ALBRECHT	88h	ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> COAN 00 obtains a fitted signal yield of  $15.9^{+5.7}_{-5.2}$  events. A search for contamination by  $K^*(1410)$  yielded a rate consistent with 0; the central value assumes no contamination.

### $\Gamma(K_2(1770)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{70}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.2 × 10<sup>-3</sup></b>	90	<sup>1</sup> LESIAK	92	CBAL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> LESIAK 92 set a limit on the inclusive process  $B(b \rightarrow s\gamma) < 2.8 \times 10^{-3}$  at 90% CL for the range of masses of 892–2045 MeV, independent of assumptions about s-quark hadronization.

### $\Gamma(K_3^*(1780)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{71}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.7 × 10<sup>-5</sup></b>	90	<sup>1</sup> NISHIDA	05	BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0 × 10 <sup>-3</sup>	90	ALBRECHT	88h	ARG $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses $B(K_3^*(1780) \rightarrow \eta K) = 0.11^{+0.05}_{-0.04}$ .				

### $\Gamma(K_4^*(2045)\gamma)/\Gamma_{\text{total}}$ $\Gamma_{72}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0 × 10<sup>-3</sup></b>	90	<sup>1</sup> LESIAK	92	CBAL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> LESIAK 92 set a limit on the inclusive process  $B(b \rightarrow s\gamma) < 2.8 \times 10^{-3}$  at 90% CL for the range of masses of 892–2045 MeV, independent of assumptions about s-quark hadronization.

### $\Gamma(K\eta(958))/\Gamma_{\text{total}}$ $\Gamma_{73}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>(8.3 ± 0.9 ± 0.7) × 10<sup>-5</sup></b> <b>0.8</b>	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(K^*(892)\eta(958))/\Gamma_{\text{total}}$ $\Gamma_{74}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>4.1 ± 1.0 ± 0.9 ± 0.5</b>		<sup>1</sup> AUBERT	07e	BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<22	90	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

### $\Gamma(K\eta)/\Gamma_{\text{total}}$ $\Gamma_{75}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.2 × 10<sup>-6</sup></b>	90	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .				

### $\Gamma(K^*(892)\eta)/\Gamma_{\text{total}}$ $\Gamma_{76}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>(1.80 ± 0.49 ± 0.18) × 10<sup>-5</sup></b> <b>0.43</b>	<sup>1</sup> RICHICHI	00	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

### $\Gamma(K\phi\phi)/\Gamma_{\text{total}}$ $\Gamma_{77}/\Gamma$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.3 ± 0.9 ± 0.3</b> <b>0.8</b>	<sup>1</sup> HUANG	03	BELL $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of charged and neutral  $B$  meson pairs and isospin symmetry.

### $\Gamma(\bar{B} \rightarrow \bar{3}\gamma)/\Gamma_{\text{total}}$ $\Gamma_{78}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.49 ± 0.19 OUR AVERAGE</b>			
3.75 ± 0.18 ± 0.35	<sup>1,2</sup> SAITO	15	BELL $e^+e^- \rightarrow \Upsilon(4S)$
3.52 ± 0.20 ± 0.51	<sup>1,3</sup> LEES	12v	BABR $e^+e^- \rightarrow \Upsilon(4S)$
3.32 ± 0.16 ± 0.31	<sup>1,4</sup> LEES	12v	BABR $e^+e^- \rightarrow \Upsilon(4S)$
3.47 ± 0.15 ± 0.40	<sup>1,5</sup> LIMOSANI	09	BELL $e^+e^- \rightarrow \Upsilon(4S)$
3.90 ± 0.91 ± 0.64	<sup>1,6</sup> AUBERT	08o	BABR $e^+e^- \rightarrow \Upsilon(4S)$
3.29 ± 0.44 ± 0.29	<sup>1,7</sup> CHEN	01c	CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.30 ± 0.08 ± 0.30	<sup>8</sup> DEL-AMO-SA.	10M	BABR $e^+e^- \rightarrow \Upsilon(4S)$
4.3 ± 0.3 ± 0.7	<sup>9</sup> AUBERT	09u	BABR Repl. by DEL-AMO-SANCHEZ 10M
3.92 ± 0.31 ± 0.47	<sup>1,10</sup> AUBERT,BE	06B	BABR Repl. by LEES 12v
3.49 ± 0.20 ± 0.59 ± 0.46	<sup>1,11</sup> AUBERT,B	05R	BABR Repl. by LEES 12u
3.50 ± 0.32 ± 0.31	<sup>1,12</sup> KOPPENBURG	04	BELL Repl. by LIMOSANI 09
3.36 ± 0.53 ± 0.65 ± 0.68	<sup>13</sup> ABE	01F	BELL Repl. by SAITO 15
2.32 ± 0.57 ± 0.35	ALAM	95	CLE2 Repl. by CHEN 01c

<sup>1</sup> We extrapolate the measured value to  $E_\gamma > 1.6$  GeV using the method of BUCH-MUELLER 06 (average of three theoretical models).

<sup>2</sup> SAITO 15 measured  $(3.51 \pm 0.17 \pm 0.33) \times 10^{-4}$  using a sum-of-exclusive approach in which 38 of the hadronic final states with  $m_{X_c} < 2.8$  GeV/ $c^2$  are reconstructed. The cut of minimum photon energy is  $E_\gamma > 1.9$  GeV.

<sup>3</sup> Reports  $(3.29 \pm 0.19 \pm 0.48) \times 10^{-4}$  for  $E_\gamma > 1.9$  GeV.

<sup>4</sup> Reports  $(3.21 \pm 0.15 \pm 0.29 \pm 0.08) \times 10^{-4}$  for  $1.8 < E_\gamma < 2.8$  GeV, where the last systematic uncertainty is for model dependency. Results with other cutoffs are also reported.

<sup>5</sup> The measurement reported is  $(3.45 \pm 0.15 \pm 0.40) \times 10^{-4}$  for  $E_\gamma > 1.7$  GeV.

<sup>6</sup> Uses a fully reconstructed  $B$  meson as a tag on the recoil side. The measurement reported is  $(3.66 \pm 0.85 \pm 0.60) \times 10^{-4}$  for  $E_\gamma > 1.9$  GeV.

<sup>7</sup> The measurement reported is  $(3.21 \pm 0.43^{+0.32}_{-0.29}) \times 10^{-4}$  for  $E_\gamma > 2.0$  GeV.

<sup>8</sup> Measured using sums of seven exclusive final states  $B \rightarrow X_{d(s)}\gamma$  where  $X_{d(s)}$  is a nonstrange (strange) charmless hadronic system in mass range 0.5–2.0 GeV/ $c^2$ .

<sup>9</sup> Measured using sums of seven exclusive final states  $B \rightarrow X_{d(s)}\gamma$  where  $X_{d(s)}$  is a nonstrange (strange) charmless hadronic system in mass range 0.6–1.8 GeV/ $c^2$ .

<sup>10</sup> The measurement reported is  $(3.67 \pm 0.29 \pm 0.45) \times 10^{-4}$  for  $E_\gamma > 1.9$  GeV.

See key on page 601

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

- <sup>11</sup> The measurement reported is  $(3.27 \pm 0.18^{+0.55}_{-0.42}) \times 10^{-4}$  for  $E_\gamma > 1.9$  GeV.  
<sup>12</sup> The measurement reported is  $(3.55 \pm 0.32 \pm 0.32) \times 10^{-4}$  for  $E_\gamma > 1.8$  GeV.  
<sup>13</sup> ABE 01F reports their systematic errors  $(\pm 0.42^{+0.50}_{-0.54}) \times 10^{-4}$ , where the second error is due to the theoretical uncertainty. We combine them in quadrature.

$\Gamma(\bar{b} \rightarrow \bar{s}\gamma)/\Gamma_{\text{total}}$	$\Gamma_{79}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID TECN COMMENT
<b><math>9.2 \pm 2.0 \pm 2.3</math></b>	<sup>1</sup> DEL-AMO-SA...10M BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$14 \pm 5 \pm 4$	<sup>2</sup> AUBERT 09u BABR Repl. by DEL-AMO-SANCHEZ 10M
<sup>1</sup> Measured using sums of seven exclusive final states $B \rightarrow X_{d(s)}\gamma$ where $X_{d(s)}$ is a nonstrange (strange) charmless hadronic system in mass range 0.5–2.0 GeV/c <sup>2</sup> .	
<sup>2</sup> Measured using sums of seven exclusive final states $B \rightarrow X_{d(s)}\gamma$ where $X_{d(s)}$ is a nonstrange (strange) charmless hadronic system in mass range 0.6–1.8 GeV/c <sup>2</sup> .	

$\Gamma(\bar{b} \rightarrow \bar{s}\gamma)/\Gamma(\bar{b} \rightarrow \bar{s}\gamma)$	$\Gamma_{79}/\Gamma_{78}$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.040 \pm 0.009 \pm 0.010</math></b>	<sup>1</sup> DEL-AMO-SA...10M BABR $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.033 \pm 0.013 \pm 0.009$	<sup>2</sup> AUBERT 09u BABR Repl. by DEL-AMO-SANCHEZ 10M
<sup>1</sup> Measured using sums of seven exclusive final states $B \rightarrow X_{d(s)}\gamma$ where $X_{d(s)}$ is a nonstrange (strange) charmless hadronic system in mass range 0.5–2.0 GeV/c <sup>2</sup> .	
<sup>2</sup> Measured using sums of seven exclusive final states $B \rightarrow X_{d(s)}\gamma$ where $X_{d(s)}$ is a nonstrange (strange) charmless hadronic system in mass range 0.6–1.8 GeV/c <sup>2</sup> .	

$\Gamma(\bar{b} \rightarrow \bar{s}\text{gluon})/\Gamma_{\text{total}}$	$\Gamma_{80}/\Gamma$
VALUE CL% EVTS	DOCUMENT ID TECN COMMENT
<b><math>&lt;0.068</math></b>	90 <sup>1</sup> COAN 98 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<0.08$	2 <sup>2</sup> ALBRECHT 95D ARG $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> COAN 98 uses $D$ - $\ell$ correlation.	
<sup>2</sup> ALBRECHT 95D use full reconstruction of one $B$ decay as tag. Two candidate events for charmless $B$ decay can be interpreted as either $b \rightarrow \text{sgluon}$ or $b \rightarrow u$ transition. If interpreted as $b \rightarrow \text{sgluon}$ they find a branching ratio of $\sim 0.026$ or the upper limit quoted above. Result is highly model dependent.	

$\Gamma(\eta \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{81}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>2.61 \pm 0.30^{+0.44}_{-0.74}</math></b>	<sup>1</sup> NISHIMURA 10 BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$1.69 \pm 0.29^{+0.36}_{-0.62}$	<sup>2</sup> NISHIMURA 10 BELL $e^+e^- \rightarrow \Upsilon(4S)$
$<4.4$	90 <sup>3</sup> BROWDER 98 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Uses $B \rightarrow \eta X_S$ with $0.4 < m_{X_S} < 2.6$ GeV/c <sup>2</sup> .	
<sup>2</sup> Uses $B \rightarrow \eta X_S$ with $1.8 < m_{X_S} < 2.6$ GeV/c <sup>2</sup> .	
<sup>3</sup> BROWDER 98 search for high momentum $B \rightarrow \eta X_S$ between 2.1 and 2.7 GeV/c.	

$\Gamma(\eta' \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{82}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>4.2 \pm 0.9</math> OUR AVERAGE</b>	
$3.9 \pm 0.8 \pm 0.9$	<sup>1</sup> AUBERT,B 04F BABR $e^+e^- \rightarrow \Upsilon(4S)$
$4.6 \pm 1.1 \pm 0.6$	<sup>2</sup> BONVICINI 03 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$6.2 \pm 1.6^{+1.3}_{-2.0}$	<sup>3</sup> BROWDER 98 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> AUBERT,B 04F reports branching ratio $B \rightarrow \eta' X_S$ for high momentum $\eta'$ between 2.0 and 2.7 GeV/c in the $\Upsilon(4S)$ center-of-mass frame. $X_S$ represents a recoil system consisting of a kaon and zero to four pions.	
<sup>2</sup> BONVICINI 03 observed a signal of $61.2 \pm 13.9$ events in $B \rightarrow \eta' X_{nc}$ production for high momentum $\eta'$ between 2.0 and 2.7 GeV/c in the $\Upsilon(4S)$ center-of-mass frame. The $X_{nc}$ denotes “charmless” hadronic states recoiling against $\eta'$ . The second error combines systematic and background subtraction uncertainties in quadrature.	
<sup>3</sup> BROWDER 98 observed a signal of $39.0 \pm 11.6$ events in high momentum $B \rightarrow \eta' X_S$ production between 2.0 and 2.7 GeV/c. The branching fraction is based on the interpretation of $b \rightarrow \text{sg}$ , where the last error includes additional uncertainties due to the color-suppressed $b \rightarrow \text{backgrounds}$ .	

$\Gamma(K^+ \text{ gluon (charmless)})/\Gamma_{\text{total}}$	$\Gamma_{83}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>&lt;1.87</math></b>	90 <sup>1</sup> DEL-AMO-SA...11 BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> $B \rightarrow K^+ X$ with $m_X < 1.69$ GeV/c <sup>2</sup> .	

$\Gamma(K^0 \text{ gluon (charmless)})/\Gamma_{\text{total}}$	$\Gamma_{84}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>1.95^{+0.51}_{-0.45} \pm 0.50</math></b>	<sup>1</sup> DEL-AMO-SA...11 BABR $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> $B \rightarrow K^0 X$ with $m_X < 1.69$ GeV/c <sup>2</sup> .	

$\Gamma(\rho\gamma)/\Gamma_{\text{total}}$	$\Gamma_{85}/\Gamma$
VALUE (units $10^{-6}$ )	CL% DOCUMENT ID TECN COMMENT
<b><math>1.39 \pm 0.25</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.
$1.73^{+0.34}_{-0.32} \pm 0.17$	<sup>1,2</sup> AUBERT 08BH BABR $e^+e^- \rightarrow \Upsilon(4S)$
$1.21^{+0.24}_{-0.22} \pm 0.12$	<sup>1,2</sup> TANIGUCHI 08 BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$1.36^{+0.29}_{-0.27} \pm 0.10$	<sup>1,3</sup> AUBERT 07L BABR Repl. by AUBERT 08BH
$<1.9$	90 <sup>1,3</sup> AUBERT 04c BABR Repl. by AUBERT 07L
$<14$	90 <sup>1,4</sup> COAN 00 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	
<sup>2</sup> Assumes $\Gamma(B \rightarrow \rho\gamma) = \Gamma(B^+ \rightarrow \rho^+\gamma) = 2 \Gamma(B^0 \rightarrow \rho^0\gamma)$ and uses lifetime ratio of $\tau_{B^+}/\tau_{B^0} = 1.071 \pm 0.009$ .	
<sup>3</sup> Assumes $\Gamma(B \rightarrow \rho\gamma) = \Gamma(B^+ \rightarrow \rho^+\gamma) = 2 \Gamma(B^0 \rightarrow \rho^0\gamma)$ and uses lifetime ratio of $\tau_{B^+}/\tau_{B^0} = 1.083 \pm 0.017$ .	
<sup>4</sup> COAN 00 reports $B(B \rightarrow \rho\gamma)/B(B \rightarrow K^*(892)\gamma) < 0.32$ at 90%CL and scaled by the central value of $B(B \rightarrow K^*(892)\gamma) = (4.24 \pm 0.54 \pm 0.32) \times 10^{-5}$ .	

$\Gamma(\rho\gamma)/\Gamma(K^*(892)\gamma)$	$\Gamma_{85}/\Gamma_{66}$
VALUE (units $10^{-2}$ )	DOCUMENT ID TECN COMMENT
<b><math>3.02^{+0.60+0.26}_{-0.55-0.28}</math></b>	TANIGUCHI 08 BELL $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\rho/\omega\gamma)/\Gamma_{\text{total}}$	$\Gamma_{86}/\Gamma$
VALUE (units $10^{-6}$ )	CL% DOCUMENT ID TECN COMMENT
<b><math>1.30 \pm 0.23</math> OUR AVERAGE</b>	Error includes scale factor of 1.2.
$1.63^{+0.30}_{-0.28} \pm 0.16$	<sup>1,2,3</sup> AUBERT 08BH BABR $e^+e^- \rightarrow \Upsilon(4S)$
$1.14 \pm 0.20^{+0.10}_{-0.12}$	<sup>1,3</sup> TANIGUCHI 08 BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$1.25^{+0.25}_{-0.24} \pm 0.09$	<sup>4</sup> AUBERT 07L BABR Repl. by AUBERT 08BH
$1.32^{+0.34+0.10}_{-0.31-0.09}$	<sup>4</sup> MOHAPATRA 06 BELL Repl. by TANIGUCHI 08
$0.6 \pm 0.3 \pm 0.1$	<sup>4</sup> AUBERT 05 BABR Repl. by AUBERT 07L
$<1.4$	<sup>4</sup> MOHAPATRA 05 BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Assumes $\Gamma(B \rightarrow \rho\gamma) = \Gamma(B^+ \rightarrow \rho^+\gamma) = 2 \Gamma(B^0 \rightarrow \rho^0\gamma)$ and uses lifetime ratio of $\tau_{B^+}/\tau_{B^0} = 1.071 \pm 0.009$ .	
<sup>2</sup> Also reports $ V_{td}/V_{ts}  = 0.233^{+0.025+0.022}_{-0.024-0.021}$ .	
<sup>3</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ .	
<sup>4</sup> Assumes $\Gamma(B \rightarrow \rho\gamma) = \Gamma(B^+ \rightarrow \rho^+\gamma) = 2 \Gamma(B^0 \rightarrow \rho^0\gamma)$ and uses lifetime ratio of $\tau_{B^+}/\tau_{B^0} = 1.083 \pm 0.017$ .	

$\Gamma(\rho/\omega\gamma)/\Gamma(K^*(892)\gamma)$	$\Gamma_{86}/\Gamma_{66}$
VALUE (units $10^{-2}$ )	CL% DOCUMENT ID TECN COMMENT
<b><math>2.84 \pm 0.50^{+0.27}_{-0.29}</math></b>	<sup>1</sup> TANIGUCHI 08 BELL $e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<3.5$	90 MOHAPATRA 05 BELL Repl. by TANIGUCHI 08
<sup>1</sup> Also reports $ V_{td}/V_{ts}  = 0.195^{+0.020}_{-0.019} \pm 0.015$ .	

$\Gamma(\pi^\pm \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{87}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>3.585 \pm 0.025 \pm 0.070</math></b>	<sup>1</sup> ALBRECHT 93i ARG $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> ALBRECHT 93 excludes $\pi^\pm$ from $K_S^0$ and $\Lambda$ decays. If included, they find $4.105 \pm 0.025 \pm 0.080$ .	

$\Gamma(\pi^0 \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{88}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>2.35 \pm 0.02 \pm 0.11</math></b>	<sup>1</sup> ABE 01j BELL $e^+e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> From fully inclusive $\pi^0$ yield with no corrections from decays of $K_S^0$ or other particles.	

$\Gamma(\eta \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{89}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.176 \pm 0.011 \pm 0.012</math></b>	KUBOTA 96 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\rho^0 \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{90}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.208 \pm 0.042 \pm 0.032</math></b>	ALBRECHT 94j ARG $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\omega \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{91}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>&lt;0.81</math></b>	90 ALBRECHT 94j ARG $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\phi \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{92}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.0343 \pm 0.0012</math> OUR AVERAGE</b>	
$0.0353 \pm 0.0005 \pm 0.0030$	HUANG 07 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
$0.0341 \pm 0.0006 \pm 0.0012$	AUBERT 04s BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.0390 \pm 0.0030 \pm 0.0035$	ALBRECHT 94j ARG $e^+e^- \rightarrow \Upsilon(4S)$
$0.023 \pm 0.006 \pm 0.005$	BORTOLETTO86 CLEO $e^+e^- \rightarrow \Upsilon(4S)$

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

$\Gamma(\phi K^*(892))/\Gamma_{\text{total}}$	$\Gamma_{93}/\Gamma$
VALUE	CL%
$<2.2 \times 10^{-5}$	90
1 BERGFELD 98 CLE2	

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .

$\Gamma(\pi^+ \text{ gluon (charmless)})/\Gamma_{\text{total}}$	$\Gamma_{95}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID
$3.72^{+0.50}_{-0.47} \pm 0.59$	1 DEL-AMO-SA..11 BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup>  $B \rightarrow \pi^+ X$  with  $m_X < 1.71 \text{ GeV}/c^2$ .

$\Gamma(\Lambda_c^+ / \bar{\Lambda}_c^- \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{96}/\Gamma$
VALUE (%)	CL%
$3.54 \pm 0.32^{+0.19}_{-0.18}$	1 AUBERT 07c BABR $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

6.4 $\pm 0.8 \pm 0.8$	2 CRAWFORD 92 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
14 $\pm 9$	3 ALBRECHT 88e ARG $e^+e^- \rightarrow \Upsilon(4S)$
$<11.2$	4 ALAM 87 CLEO $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AUBERT 07c reports  $0.045 \pm 0.003 \pm 0.012$  from a measurement of  $[\Gamma(B \rightarrow \Lambda_c^+ / \bar{\Lambda}_c^- \text{ anything})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (5.0 \pm 1.3) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> CRAWFORD 92 result derived from lepton baryon correlations. Assumes all charmed baryons in  $B^0$  and  $B^\pm$  decay are  $\Lambda_c$ .

<sup>3</sup> ALBRECHT 88e measured  $\text{B}(B \rightarrow \Lambda_c^+ X) \cdot \text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (0.30 \pm 0.12 \pm 0.06)\%$  and used  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (2.2 \pm 1.0)\%$  from ABRAMS 80 to obtain above number.

<sup>4</sup> Assuming all baryons result from charmed baryons, ALAM 86 conclude the branching fraction is  $7.4 \pm 2.9\%$ . The limit given above is model independent.

$\Gamma(\Lambda_c^+ \text{ anything})/\Gamma(\bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{97}/\Gamma_{98}$
VALUE	DOCUMENT ID
$0.19 \pm 0.13 \pm 0.04$	1 AMMAR 97 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AMMAR 97 uses a high-momentum lepton tag ( $P_\ell > 1.4 \text{ GeV}/c^2$ ).

$\Gamma(\bar{\Lambda}_c^- \mu^+ \text{ anything})/\Gamma(\bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{101}/\Gamma_{98}$
VALUE (units $10^{-2}$ )	DOCUMENT ID
$-2.0 \pm 2.0 \pm 1.9$	LEES 12 BABR $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\bar{\Lambda}_c^- \ell^+ \text{ anything})/\Gamma(\Lambda_c^+ / \bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{99}/\Gamma_{96}$
VALUE	CL%
$<2.5 \times 10^{-2}$	90
1 LEES 12 BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> LEES 12 quotes also the measurement  $\Gamma(B \rightarrow \bar{\Lambda}_c^- \ell^+ \text{ anything})/\Gamma(B \rightarrow \Lambda_c^+ / \bar{\Lambda}_c^- \text{ anything}) = (1.2 \pm 0.7 \pm 0.4) \times 10^{-2}$ .

$\Gamma(\bar{\Lambda}_c^- e^+ \text{ anything})/\Gamma(\Lambda_c^+ / \bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{100}/\Gamma_{96}$
VALUE	CL%
$<0.05$	90
1 BONVICINI 98 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BONVICINI 98 uses the electron with momentum above  $0.6 \text{ GeV}/c$ .

$\Gamma(\bar{\Lambda}_c^- e^+ \text{ anything})/\Gamma(\bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{100}/\Gamma_{98}$
VALUE (units $10^{-2}$ )	DOCUMENT ID
$2.5 \pm 1.1 \pm 0.6$	1 LEES 12 BABR $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses the full reconstruction of the recoiling  $B$  in a hadronic decay as a tag.

$\Gamma(\bar{\Lambda}_c^- \ell^+ \text{ anything})/\Gamma(\bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{99}/\Gamma_{98}$
VALUE	CL%
$<3.5 \times 10^{-2}$	90
1 LEES 12 BABR $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> LEES 12 quotes also the measurement  $\Gamma(B \rightarrow \bar{\Lambda}_c^- \ell^+ \text{ anything})/\Gamma(B \rightarrow \bar{\Lambda}_c^- \text{ anything}) = (1.7 \pm 1.0 \pm 0.6) \times 10^{-2}$ .

$\Gamma(\bar{\Lambda}_c^- p \text{ anything})/\Gamma(\Lambda_c^+ / \bar{\Lambda}_c^- \text{ anything})$	$\Gamma_{102}/\Gamma_{96}$
VALUE	DOCUMENT ID
$0.57 \pm 0.05 \pm 0.05$	BONVICINI 98 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

$\Gamma(\bar{\Lambda}_c^- p e^+ \nu_e)/\Gamma(\bar{\Lambda}_c^- p \text{ anything})$	$\Gamma_{103}/\Gamma_{102}$
VALUE	CL%
$<0.04$	90
1 BONVICINI 98 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> BONVICINI 98 uses the electron with momentum above  $0.6 \text{ GeV}/c$ .

$\Gamma(\bar{\Sigma}_c^- \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{104}/\Gamma$
VALUE	EVTS
$0.0033 \pm 0.0017 \pm 0.0002$	77
1 PROCARIO 94 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> PROCARIO 94 reports  $[\Gamma(B \rightarrow \bar{\Sigma}_c^- \text{ anything})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)] = 0.00021 \pm 0.00008 \pm 0.00007$  which we divide by our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{\Sigma}_c^- \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{105}/\Gamma$
VALUE	CL%
$<8 \times 10^{-3}$	90
1 PROCARIO 94 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> PROCARIO 94 reports  $[\Gamma(B \rightarrow \bar{\Sigma}_c^- \text{ anything})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)] < 0.00048$  which we divide by our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 6.35 \times 10^{-2}$ .

$\Gamma(\bar{\Sigma}_c^0 \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{106}/\Gamma$
VALUE	EVTS
$0.0036 \pm 0.0017 \pm 0.0002$	76
1 PROCARIO 94 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> PROCARIO 94 reports  $[\Gamma(B \rightarrow \bar{\Sigma}_c^0 \text{ anything})/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)] = 0.00023 \pm 0.00008 \pm 0.00007$  which we divide by our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{\Sigma}_c^0 N(N=p \text{ or } n))/\Gamma_{\text{total}}$	$\Gamma_{107}/\Gamma$
VALUE	CL%
$<1.2 \times 10^{-3}$	90
1 PROCARIO 94 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> PROCARIO 94 reports  $< 0.0017$  from a measurement of  $[\Gamma(B \rightarrow \bar{\Sigma}_c^0 N(N=p \text{ or } n))/\Gamma_{\text{total}}] \times [\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)]$  assuming  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 0.043$ , which we rescale to our best value  $\text{B}(\Lambda_c^+ \rightarrow p K^- \pi^+) = 6.35 \times 10^{-2}$ .

$\Gamma(\Xi_c^0 \text{ anything}, \Xi_c^0 \rightarrow \Xi^- \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{108}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID
<b><math>0.193 \pm 0.030</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.
$0.211 \pm 0.019 \pm 0.025$	1 AUBERT,B 05M BABR $e^+e^- \rightarrow \Upsilon(4S)$
$0.144 \pm 0.048 \pm 0.021$	2 BARISH 97 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> The yield is obtained by requiring the momentum  $P < 2.15 \text{ GeV}/c$ .

<sup>2</sup> BARISH 97 find  $79 \pm 27 \Xi_c^0$  events.

$\Gamma(\Xi_c^+, \Xi_c^+ \rightarrow \Xi^- \pi^+ \pi^+)/\Gamma_{\text{total}}$	$\Gamma_{109}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID
<b><math>0.453 \pm 0.096^{+0.085}_{-0.065}</math></b>	1 BARISH 97 CLE2 $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> BARISH 97 find  $125 \pm 28 \Xi_c^+$  events.

$\Gamma(p/\bar{p} \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{110}/\Gamma$
VALUE	EVTS
<b><math>0.080 \pm 0.004</math> OUR AVERAGE</b>	
$0.080 \pm 0.005 \pm 0.005$	ALBRECHT 93i ARG $e^+e^- \rightarrow \Upsilon(4S)$
$0.080 \pm 0.005 \pm 0.003$	CRAWFORD 92 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
$0.082 \pm 0.005^{+0.013}_{-0.010}$	2163 1 ALBRECHT 89k ARG $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$>0.021$  2 ALAM 83b CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 89k include direct and nondirect protons.

<sup>2</sup> ALAM 83b reported their result as  $> 0.036 \pm 0.006 \pm 0.009$ . Data are consistent with equal yields of  $p$  and  $\bar{p}$ . Using assumed yields below cut,  $\text{B}(B \rightarrow p + X) = 0.03$  not including protons from  $\Lambda$  decays.

$\Gamma(p/\bar{p} \text{ (direct) anything})/\Gamma_{\text{total}}$	$\Gamma_{111}/\Gamma$
VALUE	EVTS
<b><math>0.055 \pm 0.005</math> OUR AVERAGE</b>	
$0.055 \pm 0.005 \pm 0.0035$	ALBRECHT 93i ARG $e^+e^- \rightarrow \Upsilon(4S)$
$0.056 \pm 0.006 \pm 0.005$	CRAWFORD 92 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
$0.055 \pm 0.016$	1220 1 ALBRECHT 89k ARG $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ALBRECHT 89k subtract contribution of  $\Lambda$  decay from the inclusive proton yield.

$\Gamma(\bar{p} e^+ \nu_e \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{112}/\Gamma$
VALUE	CL%
$< 5.9 \times 10^{-4}$	90
1 ADAM 03b CLE2 $e^+e^- \rightarrow \Upsilon(4S)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<16 \times 10^{-4}$	90
ALBRECHT 90H ARG $e^+e^- \rightarrow \Upsilon(4S)$	

<sup>1</sup> Based on  $V-A$  model.

$\Gamma(\Lambda/\bar{\Lambda} \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{113}/\Gamma$
VALUE	EVTS
<b><math>0.040 \pm 0.005</math> OUR AVERAGE</b>	
$0.038 \pm 0.004 \pm 0.006$	2998 CRAWFORD 92 CLEO $e^+e^- \rightarrow \Upsilon(4S)$
$0.042 \pm 0.005 \pm 0.006$	943 ALBRECHT 89k ARG $e^+e^- \rightarrow \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.022 \pm 0.003 \pm 0.0022$  1 ACKERSTAFF 97N OPAL  $e^+e^- \rightarrow Z$

$>0.011$  2 ALAM 83b CLEO  $e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> ACKERSTAFF 97N assumes  $\text{B}(b \rightarrow B) = 0.868 \pm 0.041$ , i.e., an admixture of  $B^0, B^\pm$ , and  $B_s$ .

<sup>2</sup> ALAM 83b reported their result as  $> 0.022 \pm 0.007 \pm 0.004$ . Values are for  $(\text{B}(\Lambda X) + \text{B}(\bar{\Lambda} X))/2$ . Data are consistent with equal yields of  $p$  and  $\bar{p}$ . Using assumed yields below cut,  $\text{B}(B \rightarrow \Lambda X) = 0.03$ .

See key on page 601

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

$\Gamma(\Lambda \text{ anything})/\Gamma(\bar{\Lambda} \text{ anything})$	$\Gamma_{114}/\Gamma_{115}$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.43 \pm 0.09 \pm 0.07</math></b>	<sup>1</sup> AMMAR 97 CLE2 $e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> AMMAR 97 uses a high-momentum lepton tag ( $P_\ell > 1.4 \text{ GeV}/c^2$ ).

$\Gamma(\Xi^-/\Xi^+ \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{116}/\Gamma$
VALUE EVTS	DOCUMENT ID TECN COMMENT
<b><math>0.0027 \pm 0.0006</math> OUR AVERAGE</b>	
$0.0027 \pm 0.0005 \pm 0.0004$	147 CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
$0.0028 \pm 0.0014$	54 ALBRECHT 89K ARG $e^+e^- \rightarrow \gamma(4S)$

$\Gamma(\text{baryons anything})/\Gamma_{\text{total}}$	$\Gamma_{117}/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.068 \pm 0.005 \pm 0.003</math></b>	<sup>1</sup> ALBRECHT 92o ARG $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.076 \pm 0.014$	<sup>2</sup> ALBRECHT 89K ARG $e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> ALBRECHT 92o result is from simultaneous analysis of  $p$  and  $\Lambda$  yields,  $p\bar{p}$  and  $\Lambda\bar{\Lambda}$  correlations, and various lepton-baryon and lepton-baryon-antibaryon correlations. Supersedes ALBRECHT 89K.

<sup>2</sup> ALBRECHT 89K obtain this result by adding their their measurements ( $5.5 \pm 1.6$ )% for direct protons and ( $4.2 \pm 0.5 \pm 0.6$ )% for inclusive  $\Lambda$  production. They then assume ( $5.5 \pm 1.6$ )% for neutron production and add it in also. Since each  $B$  decay has two baryons, they divide by 2 to obtain ( $7.6 \pm 1.4$ )%.

$\Gamma(p\bar{p} \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{118}/\Gamma$
Includes $p$ and $\bar{p}$ from $\Lambda$ and $\bar{\Lambda}$ decay.	
VALUE EVTS	DOCUMENT ID TECN COMMENT
<b><math>0.0247 \pm 0.0023</math> OUR AVERAGE</b>	
$0.024 \pm 0.001 \pm 0.004$	92 CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
$0.025 \pm 0.002 \pm 0.002$	918 ALBRECHT 89K ARG $e^+e^- \rightarrow \gamma(4S)$

$\Gamma(p\bar{p} \text{ anything})/\Gamma(p/\bar{p} \text{ anything})$	$\Gamma_{118}/\Gamma_{110}$
Includes $p$ and $\bar{p}$ from $\Lambda$ and $\bar{\Lambda}$ decay.	
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.30 \pm 0.02 \pm 0.05$	<sup>1</sup> CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> CRAWFORD 92 value is not independent of their $\Gamma(p\bar{p} \text{ anything})/\Gamma_{\text{total}}$ value.	

$\Gamma(\Lambda\bar{\Lambda}/\Lambda p \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{119}/\Gamma$
Includes $p$ and $\bar{p}$ from $\Lambda$ and $\bar{\Lambda}$ decay.	
VALUE EVTS	DOCUMENT ID TECN COMMENT
<b><math>0.025 \pm 0.004</math> OUR AVERAGE</b>	
$0.029 \pm 0.005 \pm 0.005$	92 CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
$0.023 \pm 0.004 \pm 0.003$	165 ALBRECHT 89K ARG $e^+e^- \rightarrow \gamma(4S)$

$\Gamma(\Lambda\bar{\Lambda}/\Lambda p \text{ anything})/\Gamma(\Lambda/\bar{\Lambda} \text{ anything})$	$\Gamma_{119}/\Gamma_{113}$
Includes $p$ and $\bar{p}$ from $\Lambda$ and $\bar{\Lambda}$ decay.	
VALUE	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.76 \pm 0.11 \pm 0.08$	<sup>1</sup> CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> CRAWFORD 92 value is not independent of their $[\Gamma(\Lambda\bar{\Lambda} \text{ anything}) + \Gamma(\Lambda p \text{ anything})]/\Gamma_{\text{total}}$ value.	

$\Gamma(\Lambda\bar{\Lambda} \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{120}/\Gamma$
VALUE CL% EVTS	DOCUMENT ID TECN COMMENT
<b><math>&lt;0.005</math></b>	90 CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<0.0088$	90 12 ALBRECHT 89K ARG $e^+e^- \rightarrow \gamma(4S)$

$\Gamma(\Lambda\bar{\Lambda} \text{ anything})/\Gamma(\Lambda/\bar{\Lambda} \text{ anything})$	$\Gamma_{120}/\Gamma_{113}$
VALUE CL%	DOCUMENT ID TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<0.13$	90 <sup>1</sup> CRAWFORD 92 CLEO $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> CRAWFORD 92 value is not independent of their $\Gamma(\Lambda\bar{\Lambda} \text{ anything})/\Gamma_{\text{total}}$ value.	

$\Gamma(se^+e^-)/\Gamma_{\text{total}}$	$\Gamma_{121}/\Gamma$
Test for $\Delta B = 1$ weak neutral current. Allowed by higher-order electroweak interactions.	
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT
<b><math>6.7 \pm 1.7</math> OUR AVERAGE</b>	Error includes scale factor of 2.0.
$7.69^{+0.82+0.71}_{-0.77-0.60}$	<sup>1</sup> LEES 14D BABR $e^+e^- \rightarrow \gamma(4S)$
$4.04 \pm 1.30^{+0.87}_{-0.83}$	<sup>2</sup> IWASAKI 05 BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$6.0 \pm 1.7 \pm 1.3$	<sup>2</sup> AUBERT,B 04i BABR Repl. by LEES 14D
$5.0 \pm 2.3^{+1.3}_{-1.1}$	<sup>2</sup> KANEKO 03 BELL Repl. by IWASAKI 05
$< 57$	90 GLENN 98 CLEO $e^+e^- \rightarrow \gamma(4S)$
$< 50000$	90 BEBEK 81 CLEO $e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

<sup>2</sup> Requires  $M_{\ell^+\ell^-} > 0.2 \text{ GeV}/c^2$ .

$\Gamma(s\mu^+\mu^-)/\Gamma_{\text{total}}$	$\Gamma_{122}/\Gamma$
Test for $\Delta B = 1$ weak neutral current. Allowed by higher-order electroweak interactions.	
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT
<b><math>4.3 \pm 1.0</math> OUR AVERAGE</b>	
$4.41^{+1.31+0.63}_{-1.17-0.50}$	<sup>1</sup> LEES 14D BABR $e^+e^- \rightarrow \gamma(4S)$
$4.13 \pm 1.05^{+0.85}_{-0.81}$	<sup>2</sup> IWASAKI 05 BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$5.0 \pm 2.8 \pm 1.2$	AUBERT,B 04i BABR Repl. by LEES 14D
$7.9 \pm 2.1^{+2.1}_{-1.5}$	KANEKO 03 BELL Repl. by IWASAKI 05
$< 58$	90 GLENN 98 CLEO $e^+e^- \rightarrow \gamma(4S)$
$< 17000$	90 CHADWICK 81 CLEO $e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

<sup>2</sup> Requires  $M_{\ell^+\ell^-} > 0.2 \text{ GeV}/c^2$ .

$[\Gamma(se^+e^-) + \Gamma(s\mu^+\mu^-)]/\Gamma_{\text{total}}$	$(\Gamma_{121} + \Gamma_{122})/\Gamma$
Test for $\Delta B = 1$ weak neutral current. Allowed by higher-order electroweak interactions.	
VALUE CL%	DOCUMENT ID TECN COMMENT
<b><math>&lt;4.2 \times 10^{-5}</math></b>	90 GLENN 98 CLEO $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<0.0024$	90 <sup>1</sup> BEAN 87 CLEO Repl. by GLENN 98
$<0.0062$	90 <sup>2</sup> AVERY 84 CLEO Repl. by BEAN 87
<sup>1</sup> BEAN 87 reports $[(\mu^+\mu^-) + (e^+e^-)]/2$ and we converted it.	
<sup>2</sup> Determine ratio of $B^+$ to $B^0$ semileptonic decays to be in the range 0.25–2.9.	

$\Gamma(se^+\ell^-)/\Gamma_{\text{total}}$	$\Gamma_{123}/\Gamma$
Test for $\Delta B = 1$ weak neutral current.	
VALUE (units $10^{-6}$ ) CL%	DOCUMENT ID TECN COMMENT
<b><math>5.8 \pm 1.3</math> OUR AVERAGE</b>	Error includes scale factor of 1.8.
$6.73^{+0.70+0.60}_{-0.64-0.56}$	<sup>1</sup> LEES 14D BABR $e^+e^- \rightarrow \gamma(4S)$
$4.11 \pm 0.83^{+0.85}_{-0.81}$	<sup>2</sup> IWASAKI 05 BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$5.6 \pm 1.5 \pm 1.3$	<sup>3</sup> AUBERT,B 04i BABR Repl. by LEES 14D
$6.1 \pm 1.4^{+1.4}_{-1.1}$	<sup>3</sup> KANEKO 03 BELL Repl. by IWASAKI 05
<sup>1</sup> Measured from sum of exclusive modes through $K^+$ , $K^+\pi^0$ , $K^+\pi^-$ , $K^+\pi^-\pi^0$ , $K^+\pi^-\pi^+$ , $K_S^0$ , $K_S^0\pi^0$ , $K_S^0\pi^+$ , $K_S^0\pi^+\pi^0$ , and $K_S^0\pi^+ + \pi^-$ corrected for unobserved modes.	
<sup>2</sup> Requires $M_{\ell^+\ell^-} > 0.2 \text{ GeV}/c^2$ .	
<sup>3</sup> Requires $M_{e^+e^-} > 0.2 \text{ GeV}/c^2$ .	

$\Gamma(\pi\ell^+\ell^-)/\Gamma_{\text{total}}$	$\Gamma_{124}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT
<b><math>&lt;5.9 \times 10^{-8}</math></b>	90 <sup>1</sup> LEES 13M BABR $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$<6.2 \times 10^{-8}$	90 <sup>1</sup> WEI 08A BELL $e^+e^- \rightarrow \gamma(4S)$
$<9.1 \times 10^{-8}$	90 <sup>1</sup> AUBERT 07AG BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .	

$\Gamma(\pi e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_{125}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT
<b><math>&lt;11.0 \times 10^{-8}</math></b>	90 <sup>1</sup> LEES 13M BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .	

$\Gamma(\pi\mu^+\mu^-)/\Gamma_{\text{total}}$	$\Gamma_{126}/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT
<b><math>&lt;5.0 \times 10^{-8}</math></b>	90 <sup>1</sup> LEES 13M BABR $e^+e^- \rightarrow \gamma(4S)$
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .	

$\Gamma(K e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_{127}/\Gamma$
Test for $\Delta B = 1$ weak neutral current. Allowed by higher-order electroweak interactions.	
VALUE (units $10^{-7}$ ) CL%	DOCUMENT ID TECN COMMENT
<b><math>4.4 \pm 0.6</math> OUR AVERAGE</b>	
$3.9^{+0.9}_{-0.8} \pm 0.2$	<sup>1</sup> AUBERT 09T BABR $e^+e^- \rightarrow \gamma(4S)$
$4.8^{+0.8}_{-0.7} \pm 0.3$	<sup>1</sup> WEI 09A BELL $e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$3.3^{+0.9}_{-0.8} \pm 0.2$	<sup>1</sup> AUBERT,B 06j BABR Repl. by AUBERT 09T
$7.4^{+1.8}_{-1.6} \pm 0.5$	<sup>1</sup> AUBERT 03u BABR Repl. by AUBERT,B 06j
$4.8^{+1.5}_{-1.3} \pm 0.3$	<sup>1,2</sup> ISHIKAWA 03 BELL Repl. by WEI 09A
$< 13$	90 ABE 02 BELL Repl. by ISHIKAWA 03
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\gamma(4S)$ .	
<sup>2</sup> The second error is a total of systematic uncertainties including model dependence.	

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

$\Gamma(K^*(892)e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{128}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>11.9 \pm 2.0</math> OUR AVERAGE</b>		Error includes scale factor of 1.2.		
$9.9^{+2.3}_{-2.1} \pm 0.6$		<sup>1</sup> AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$13.9^{+2.3}_{-2.0} \pm 1.2$		<sup>1</sup> WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$9.7^{+3.0}_{-2.7} \pm 1.4$		<sup>1</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$9.8^{+5.0}_{-4.2} \pm 1.1$		<sup>1</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$14.9^{+5.2+1.2}_{-4.6-1.3}$		<sup>2</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
<56	90	ABE	02 BELL	Repl. by ISHIKAWA 03

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ . The second error is a total of systematic uncertainties including model dependence.

$\Gamma(K\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{129}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>4.4 \pm 0.4</math> OUR AVERAGE</b>			
$4.2 \pm 0.4 \pm 0.2$	AALTONEN	11A1 CDF	$p\bar{p}$ at 1.96 TeV
$4.1^{+1.3}_{-1.2} \pm 0.2$	<sup>1</sup> AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$5.0 \pm 0.6 \pm 0.3$	<sup>1</sup> WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$3.5^{+1.3}_{-1.1} \pm 0.3$	<sup>1</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$4.5^{+2.3}_{-1.9} \pm 0.4$	<sup>1</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$4.8^{+1.2}_{-1.1} \pm 0.4$	<sup>1,2</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
$9.9^{+4.0+1.3}_{-3.2-1.0}$	ABE	02 BELL	Repl. by ISHIKAWA 03

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> The second error is a total of systematic uncertainties including model dependence.

$\Gamma(K\mu^+\mu^-)/\Gamma(Ke^+e^-)$   $\Gamma_{129}/\Gamma_{127}$   
VALUE DOCUMENT ID TECN COMMENT

<b><math>1.01 \pm 0.15</math> OUR AVERAGE</b>			
$1.00^{+0.31}_{-0.25} \pm 0.07$	<sup>1</sup> LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.96^{+0.44}_{-0.34} \pm 0.05$	AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.03 \pm 0.19 \pm 0.06$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.06 \pm 0.48 \pm 0.08$	AUBERT,B	06J BABR	Repl. by AUBERT 09T
<sup>1</sup> Measured in the union of $0.10 < q^2 < 8.12 \text{ GeV}^2/c^4$ and $q^2 > 10.11 \text{ GeV}^2/c^4$ . LEES 12s reports also individual measurements $\Gamma(B \rightarrow K\mu^+\mu^-)/\Gamma(B \rightarrow Ke^+e^-) = 0.74^{+0.40}_{-0.31} \pm 0.06$ for $0.10 < q^2 < 8.12 \text{ GeV}^2/c^4$ and $\Gamma(B \rightarrow K\mu^+\mu^-)/\Gamma(B \rightarrow Ke^+e^-) = 1.43^{+0.65}_{-0.44} \pm 0.12$ for $q^2 > 10.11 \text{ GeV}^2/c^4$ .			

$\Gamma(K^*(892)\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{130}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>10.6 \pm 0.9</math> OUR AVERAGE</b>				
$10.1 \pm 1.0 \pm 0.5$		AALTONEN	11A1 CDF	$p\bar{p}$ at 1.96 TeV
$13.5^{+3.5}_{-3.3} \pm 1.0$		<sup>1</sup> AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$11.0^{+1.6}_{-1.4} \pm 0.8$		<sup>1</sup> WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$8.8^{+3.5}_{-3.0} \pm 1.2$		<sup>1</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$12.7^{+7.6}_{-6.1} \pm 1.6$		<sup>1</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$11.7^{+3.6}_{-3.1} \pm 1.0$		<sup>2</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
<31	90	ABE	02 BELL	Repl. by ISHIKAWA 03

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes equal production of  $B^0$  and  $B^+$  at  $\Upsilon(4S)$ . The second error is a total of systematic uncertainties including model dependence.

$\Gamma(K^*(892)\mu^+\mu^-)/\Gamma(K^*(892)e^+e^-)$   $\Gamma_{130}/\Gamma_{128}$   
VALUE DOCUMENT ID TECN COMMENT

<b><math>0.98 \pm 0.15</math> OUR AVERAGE</b>			
$1.13^{+0.34}_{-0.26} \pm 0.10$	<sup>1</sup> LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.37^{+0.53}_{-0.40} \pm 0.09$	AUBERT	09T BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.83 \pm 0.17 \pm 0.08$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.91 \pm 0.45 \pm 0.06$	AUBERT,B	06J BABR	Repl. by AUBERT 09T

<sup>1</sup> Measured in the union of  $0.10 < q^2 < 8.12 \text{ GeV}^2/c^4$  and  $q^2 > 10.11 \text{ GeV}^2/c^4$ .  
LEES 12s reports also individual measurements  $\Gamma(B \rightarrow K^*(892)\mu^+\mu^-)/\Gamma(B \rightarrow K^*(892)e^+e^-) = 1.06^{+0.48}_{-0.33} \pm 0.08$  for  $0.10 < q^2 < 8.12 \text{ GeV}^2/c^4$  and  $\Gamma(B \rightarrow K^*(892)\mu^+\mu^-)/\Gamma(B \rightarrow K^*(892)e^+e^-) = 1.18^{+0.55}_{-0.37} \pm 0.11$  for  $q^2 > 10.11 \text{ GeV}^2/c^4$ .

$\Gamma(K\ell^+\ell^-)/\Gamma_{\text{total}}$   $\Gamma_{131}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>4.8 \pm 0.4</math> OUR AVERAGE</b>				
$4.7 \pm 0.6 \pm 0.2$		LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$4.8^{+0.5}_{-0.4} \pm 0.3$		WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.9 \pm 0.7 \pm 0.2$		<sup>1</sup> AUBERT	09T BABR	Repl. by LEES 12s
$3.4 \pm 0.7 \pm 0.2$		<sup>1</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$6.5^{+1.4}_{-1.3} \pm 0.4$		<sup>2</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$4.8^{+1.0}_{-0.9} \pm 0.3$		<sup>3</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
$7.5^{+2.5}_{-2.1} \pm 0.6$		<sup>4</sup> ABE	02 BELL	Repl. by ISHIKAWA 03
< 5.1	90	<sup>1</sup> AUBERT	02L BABR	$e^+e^- \rightarrow \Upsilon(4S)$
<17	90	<sup>5</sup> ANDERSON	01B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes all four  $B \rightarrow K\ell^+\ell^-$  modes having equal partial widths in the fit.  
<sup>3</sup> Assumes equal production rate for charge and neutral  $B$  meson pairs, isospin invariance, lepton universality for  $B \rightarrow K\ell^+\ell^-$ , and  $B(B \rightarrow K^*(892)\mu^+\mu^-) = 1.33$ . The second error is total systematic uncertainties including model dependence.  
<sup>4</sup> Assumes lepton universality.  
<sup>5</sup> The result is for di-lepton masses above 0.5 GeV.

$\Gamma(K^*(892)\ell^+\ell^-)/\Gamma_{\text{total}}$   $\Gamma_{132}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>10.5 \pm 1.0</math> OUR AVERAGE</b>				
$10.2^{+1.4}_{-1.3} \pm 0.5$		LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$10.7^{+1.1}_{-1.0} \pm 0.9$		WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$11.1^{+1.9}_{-1.8} \pm 0.7$		<sup>1</sup> AUBERT	09T BABR	Repl. by LEES 12s
$7.8^{+1.9}_{-1.7} \pm 1.1$		<sup>1</sup> AUBERT,B	06J BABR	Repl. by AUBERT 09T
$8.8^{+3.3}_{-2.9} \pm 1.0$		<sup>2</sup> AUBERT	03U BABR	Repl. by AUBERT,B 06J
$11.5^{+2.6}_{-2.4} \pm 0.8$		<sup>3</sup> ISHIKAWA	03 BELL	Repl. by WEI 09A
<31	90	<sup>1,4</sup> AUBERT	02L BABR	Repl. by AUBERT 03U
<33	90	<sup>5</sup> ANDERSON	01B CLE2	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Assumes equal production of  $B^+$  and  $B^0$  at the  $\Upsilon(4S)$ .  
<sup>2</sup> Assumes the partial width ratio of electron and muon modes to be  $\Gamma(B \rightarrow K^*(892)e^+e^-)/\Gamma(B \rightarrow K^*(892)\mu^+\mu^-) = 1.33$ .  
<sup>3</sup> Assumes equal production rate for charge and neutral  $B$  meson pairs, isospin invariance, lepton universality for  $B \rightarrow K\ell^+\ell^-$ , and  $B(B \rightarrow K^*(892)\mu^+\mu^-) = 1.33$ . The second error is total systematic uncertainties including model dependence.  
<sup>4</sup> For averaging  $K^*(892)\mu^+\mu^-$  and  $K^*(892)e^+e^-$  modes, AUBERT 02L assumed  $B(B \rightarrow K^*(892)e^+e^-)/B(B \rightarrow K^*(892)\mu^+\mu^-) = 1.2$ .  
<sup>5</sup> The result is for di-lepton masses above 0.5 GeV.

$\Gamma(K\nu\bar{\nu})/\Gamma_{\text{total}}$   $\Gamma_{133}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.7 \times 10^{-5}</math></b>	90	<sup>1,2</sup> LEES	13i BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.4 \times 10^{-5}$	90	<sup>1</sup> DEL-AMO-SA...10Q	BABR	Repl. by LEES 13i
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>2</sup> Also reported a limit $< 3.2 \times 10^{-5}$ at 90% CL obtained using a fully reconstructed hadronic $B$ -tag evnets.				

$\Gamma(K^*\nu\bar{\nu})/\Gamma_{\text{total}}$   $\Gamma_{134}/\Gamma$   
Test for  $\Delta B = 1$  weak neutral current.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;7.6 \times 10^{-5}</math></b>	90	<sup>1,2</sup> LEES	13i BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<8 \times 10^{-5}$	90	AUBERT	08Bc BABR	Repl. by LEES 13i
<sup>1</sup> Assumes equal production of $B^+$ and $B^0$ at the $\Upsilon(4S)$ . <sup>2</sup> Also reported a limit $< 7.9 \times 10^{-5}$ at 90% CL obtained using a fully reconstructed hadronic $B$ -tag evnets.				

$\Gamma(s e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{135}/\Gamma$   
Test for lepton family number conservation. Allowed by higher-order electroweak interactions.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.2 \times 10^{-5}</math></b>	90	GLENN	98 CLEO	$e^+e^- \rightarrow \Upsilon(4S)$





# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (m_{\ell\ell} > 3.2 \text{ GeV}/c^2)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.71^{+0.29}_{-0.22} \pm 0.04$	AUBERT	09N BABR	$e^+ e^- \rightarrow \Upsilon(4S)$

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (0.10 < q^2 < 0.98 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.263^{+0.045}_{-0.044} \pm 0.017$	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (1.1 < q^2 < 2.5 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.660^{+0.083}_{-0.077} \pm 0.022$	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (0.1 < q^2 < 2.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.34^{+0.09}_{-0.07}</math> OUR AVERAGE</b>			
$0.37^{+0.10+0.04}_{-0.09-0.03}$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.30 \pm 0.16 \pm 0.02$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.29^{+0.21}_{-0.18} \pm 0.02$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.60^{+0.00}_{-0.28} \pm 0.19$	<sup>1</sup> CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.00^{+0.13}_{-0.00} \pm 0.02$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.53^{+0.32}_{-0.34} \pm 0.07$	AALTONEN	11L CDF	Repl. by AALTONEN 12i
<sup>1</sup> CHATRCHYAN 13BL uses, for this bin, $1.0 < q^2 < 2.0 \text{ GeV}^2/c^4$ .			

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.77 \pm 0.05</math> OUR AVERAGE</b>			
$0.876^{+0.109}_{-0.097} \pm 0.017$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$0.80 \pm 0.08 \pm 0.06$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$0.74^{+0.10+0.02}_{-0.09-0.03}$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.65 \pm 0.17 \pm 0.03$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.37^{+0.25}_{-0.24} \pm 0.10$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.71 \pm 0.24 \pm 0.05$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.77 \pm 0.15 \pm 0.03$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.40^{+0.32}_{-0.33} \pm 0.08$	AALTONEN	11L CDF	Repl. by AALTONEN 12i
<sup>1</sup> Measured in $2.5 < q^2 < 4.0 \text{ GeV}^2/c^4$ .			

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (4.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.611^{+0.052}_{-0.053} \pm 0.017</math></b>	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (6.0 < q^2 < 8.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.579 \pm 0.046 \pm 0.015</math></b>	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (4.3 < q^2 < 8.6 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.64 \pm 0.06</math> OUR AVERAGE</b>			
$0.57 \pm 0.07 \pm 0.03$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.81^{+0.13}_{-0.12} \pm 0.05$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.68^{+0.15}_{-0.17} \pm 0.09$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.64^{+0.23}_{-0.24} \pm 0.07$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.60^{+0.06}_{-0.07} \pm 0.01$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.82^{+0.19}_{-0.23} \pm 0.07$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.448 \pm 0.033</math> OUR AVERAGE</b>			
$0.493^{+0.049}_{-0.047} \pm 0.013$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$0.39 \pm 0.05 \pm 0.04$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$0.48^{+0.08}_{-0.09} \pm 0.03$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.45^{+0.10}_{-0.11} \pm 0.04$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.47 \pm 0.14 \pm 0.03$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.17^{+0.17}_{-0.15} \pm 0.03$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.41 \pm 0.11 \pm 0.03$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.31^{+0.19}_{-0.18} \pm 0.02$	AALTONEN	11L CDF	Repl. by AALTONEN 12i
<sup>1</sup> Measured in $11.0 < q^2 < 12.5 \text{ GeV}^2/c^4$ .			

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (15.0 < q^2 < 17.0 \text{ GeV}^2/c^4)$

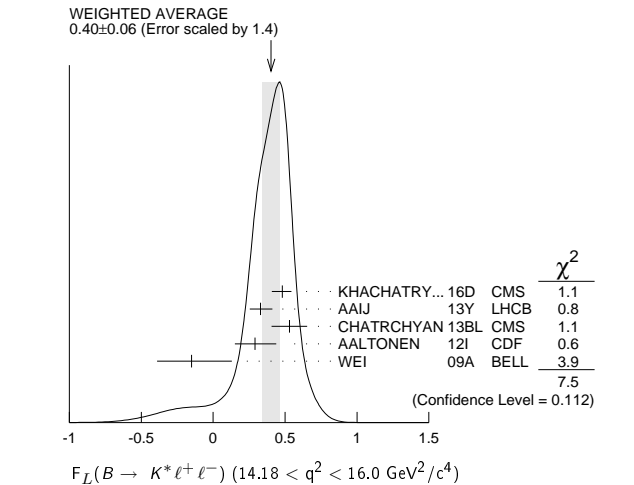
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.349 \pm 0.039 \pm 0.009</math></b>	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (17.0 < q^2 < 19.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.354^{+0.049}_{-0.048} \pm 0.025</math></b>	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (14.18 < q^2 < 16.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.40 \pm 0.06</math> OUR AVERAGE</b>	Error includes scale factor of 1.4. See the ideogram below.		
$0.48^{+0.05}_{-0.06} \pm 0.04$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$0.33^{+0.08+0.02}_{-0.07-0.03}$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.53 \pm 0.12 \pm 0.03$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.29^{+0.14}_{-0.13} \pm 0.05$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$-0.15^{+0.27}_{-0.23} \pm 0.07$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.37 \pm 0.09 \pm 0.05$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.55^{+0.17}_{-0.18} \pm 0.02$	AALTONEN	11L CDF	Repl. by AALTONEN 12i



### $F_L(B \rightarrow K^* \ell^+ \ell^-) (16.0 < q^2 < 19.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.353 \pm 0.024</math> OUR AVERAGE</b>			
$0.344^{+0.028}_{-0.030} \pm 0.008$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$0.38^{+0.05}_{-0.06} \pm 0.04$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$0.38^{+0.09}_{-0.07} \pm 0.03$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.44 \pm 0.07 \pm 0.03$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.20^{+0.19}_{-0.17} \pm 0.05$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.12^{+0.15}_{-0.13} \pm 0.02$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.26^{+0.10}_{-0.08} \pm 0.03$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.09^{+0.18}_{-0.14} \pm 0.03$	AALTONEN	11L CDF	Repl. by AALTONEN 12i
<sup>1</sup> Measured in $15.0 < q^2 < 19.0 \text{ GeV}^2/c^4$ .			

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.692 \pm 0.030</math> OUR AVERAGE</b>			
$0.690^{+0.035}_{-0.036} \pm 0.017$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$0.72 \pm 0.06$	KHACHATRY...16D	CMS	$pp$ at 7, 8 TeV
$0.65^{+0.08}_{-0.07} \pm 0.03$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.69^{+0.19}_{-0.21} \pm 0.08$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.67 \pm 0.23 \pm 0.05$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.68 \pm 0.10 \pm 0.02$	CHATRCHYAN13BL	CMS	Repl. by KHACHATRYAN 16D
$0.55 \pm 0.10 \pm 0.03$	AAIJ	12u LHCb	Repl. by AAIJ 13Y
$0.50^{+0.27}_{-0.30} \pm 0.03$	AALTONEN	11L CDF	Repl. by AALTONEN 12i
<sup>1</sup> Measured in $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ .			

See key on page 601

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

### $F_L(B \rightarrow K^* \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.33^{+0.14}_{-0.13} \pm 0.03$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.47^{+0.23}_{-0.24} \pm 0.03$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

### PARTIAL BRANCHING FRACTIONS IN $B \rightarrow K^{(*)} \ell^+ \ell^-$

#### $B(B \rightarrow K^* \ell^+ \ell^-) (q^2 < 2.0 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.68 \pm 0.23</math> OUR AVERAGE</b>			
$1.89^{+0.52}_{-0.46} \pm 0.06$	<sup>1</sup> LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.73 \pm 0.33 \pm 0.10$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$1.46^{+0.40}_{-0.35} \pm 0.11$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.98 \pm 0.40 \pm 0.09$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

<sup>1</sup> The value reported here from LEES 12s refers to  $0.1 < q^2 < 2.0 \text{ GeV}^2/c^2$ .

#### $B(B \rightarrow K^* \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.87 \pm 0.17</math> OUR AVERAGE</b>			
$0.95^{+0.35}_{-0.30} \pm 0.04$	LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.82 \pm 0.26 \pm 0.06$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$0.88^{+0.31}_{-0.27} \pm 0.07$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.00 \pm 0.38 \pm 0.09$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

#### $B(B \rightarrow K^* \ell^+ \ell^-) (4.3 < q^2 < 8.68 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.67 \pm 0.29</math> OUR AVERAGE</b>			
$1.82^{+0.56}_{-0.52} \pm 0.09$	<sup>1</sup> LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.72 \pm 0.41 \pm 0.14$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$1.37^{+0.47}_{-0.42} \pm 0.39$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.69 \pm 0.57 \pm 0.15$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

<sup>1</sup> The value reported here from LEES 12s refers to  $4.3 < q^2 < 8.12 \text{ GeV}^2/c^2$ .

#### $B(B \rightarrow K^* \ell^+ \ell^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.93 \pm 0.25</math> OUR AVERAGE</b>			
$1.86^{+0.52}_{-0.48} \pm 0.10$	<sup>1</sup> LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.77 \pm 0.34 \pm 0.11$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$2.24^{+0.44}_{-0.40} \pm 0.19$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.97 \pm 0.47 \pm 0.17$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

<sup>1</sup> The value reported here from LEES 12s refers to  $10.11 < q^2 < 12.89 \text{ GeV}^2/c^2$ .

#### $B(B \rightarrow K^* \ell^+ \ell^-) (14.18 < q^2 < 16.0 \text{ GeV}^2/c^4)$

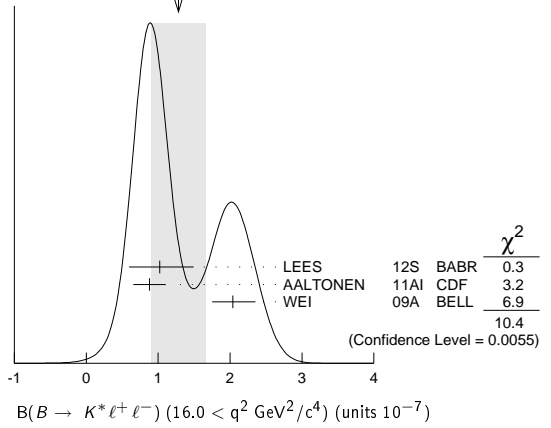
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.21 \pm 0.17</math> OUR AVERAGE</b>			
$1.46^{+0.41}_{-0.36} \pm 0.06$	<sup>1</sup> LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.21 \pm 0.24 \pm 0.07$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$1.05^{+0.29}_{-0.26} \pm 0.08$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.51 \pm 0.36 \pm 0.13$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

<sup>1</sup> The value reported here from LEES 12s refers to  $14.21 < q^2 < 16.0 \text{ GeV}^2/c^2$ .

#### $B(B \rightarrow K^* \ell^+ \ell^-) (16.0 < q^2 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.3 \pm 0.4</math> OUR AVERAGE</b>			Error includes scale factor of 2.3. See the ideogram below.
$1.02^{+0.47}_{-0.42} \pm 0.06$	LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.88 \pm 0.22 \pm 0.05$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$2.04^{+0.27}_{-0.24} \pm 0.16$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.35 \pm 0.37 \pm 0.12$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

WEIGHTED AVERAGE  
 $1.3 \pm 0.4$  (Error scaled by 2.3)



#### $B(B \rightarrow K^* \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.64 \pm 0.26</math> OUR AVERAGE</b>			
$2.05^{+0.53}_{-0.48} \pm 0.07$	LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$1.48 \pm 0.39 \pm 0.12$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$1.49^{+0.45}_{-0.40} \pm 0.12$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.60 \pm 0.54 \pm 0.14$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

#### $B(B \rightarrow K^* \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

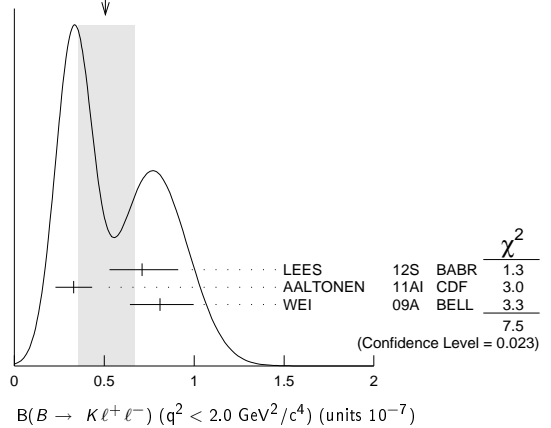
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.53 \pm 0.43 \pm 0.15</math></b>	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.98 \pm 0.55 \pm 0.18$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

#### $B(B \rightarrow K \ell^+ \ell^-) (q^2 < 2.0 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.51 \pm 0.16</math> OUR AVERAGE</b>			Error includes scale factor of 1.9. See the ideogram below.
$0.71^{+0.20}_{-0.18} \pm 0.02$	<sup>1</sup> LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.33 \pm 0.10 \pm 0.02$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$0.81^{+0.18}_{-0.16} \pm 0.05$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.38 \pm 0.16 \pm 0.03$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

<sup>1</sup> The value reported here from LEES 12s refers to  $0.1 < q^2 < 2.0 \text{ GeV}^2/c^2$ .

WEIGHTED AVERAGE  
 $0.51 \pm 0.16$  (Error scaled by 1.9)



#### $B(B \rightarrow K \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.57^{+0.10}_{-0.09} \pm 0.09</math> OUR AVERAGE</b>			Error includes scale factor of 1.2.
$0.49^{+0.15}_{-0.13} \pm 0.01$	LEES	12s BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.77 \pm 0.14 \pm 0.05$	AALTONEN	11Ai CDF	$p\bar{p}$ at 1.96 TeV
$0.46^{+0.14}_{-0.12} \pm 0.03$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.58 \pm 0.19 \pm 0.04$	AALTONEN	11L CDF	Repl. by AALTONEN 11Ai

## Meson Particle Listings

 $B^\pm/B^0$  ADMIXTURE $B(B \rightarrow K\ell^+\ell^-)$  ( $4.3 < q^2 < 8.68 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.00 \pm 0.11</math> OUR AVERAGE</b>			
$0.94^{+0.20}_{-0.19} \pm 0.02$	<sup>1</sup> LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.05 \pm 0.17 \pm 0.07$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
$1.00^{+0.19}_{-0.18} \pm 0.06$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.93 \pm 0.25 \pm 0.06$	AALTONEN	11L CDF	Repl. by AALTONEN 11AI

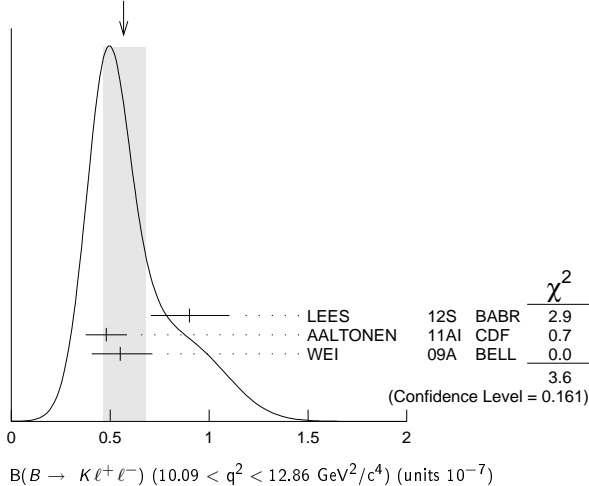
<sup>1</sup> The value reported here from LEES 12s refers to  $4.3 < q^2 < 8.12 \text{ GeV}^2/c^2$ .

 $B(B \rightarrow K\ell^+\ell^-)$  ( $10.09 < q^2 < 12.86 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.57 \pm 0.11</math> OUR AVERAGE</b>			Error includes scale factor of 1.4. See the ideogram below.
$0.90^{+0.20}_{-0.19} \pm 0.04$	<sup>1</sup> LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.48 \pm 0.10 \pm 0.03$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
$0.55^{+0.16}_{-0.14} \pm 0.03$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.72 \pm 0.17 \pm 0.05$	AALTONEN	11L CDF	Repl. by AALTONEN 11AI

<sup>1</sup> The value reported here from LEES 12s refers to  $10.11 < q^2 < 12.89 \text{ GeV}^2/c^2$ .

WEIGHTED AVERAGE  
 $0.57 \pm 0.11$  (Error scaled by 1.4)

 $B(B \rightarrow K\ell^+\ell^-)$  ( $14.18 < q^2 < 16.0 \text{ GeV}^2/c^4$ )

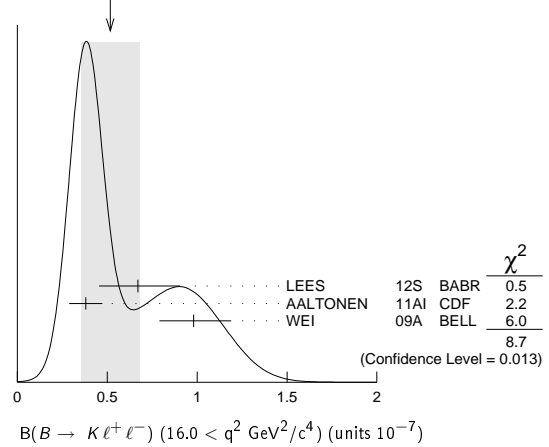
VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.49 \pm 0.07</math> OUR AVERAGE</b>			
$0.49^{+0.15}_{-0.14} \pm 0.02$	<sup>1</sup> LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.52 \pm 0.09 \pm 0.03$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
$0.38^{+0.19}_{-0.12} \pm 0.02$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.38 \pm 0.12 \pm 0.03$	AALTONEN	11L CDF	Repl. by AALTONEN 11AI

<sup>1</sup> The value reported here from LEES 12s refers to  $14.21 < q^2 < 16.0 \text{ GeV}^2/c^2$ .

 $B(B \rightarrow K\ell^+\ell^-)$  ( $16.0 < q^2 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.52 \pm 0.16</math> OUR AVERAGE</b>			Error includes scale factor of 2.1. See the ideogram below.
$0.67^{+0.23}_{-0.21} \pm 0.05$	LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$0.38 \pm 0.09 \pm 0.02$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
$0.98^{+0.20}_{-0.18} \pm 0.06$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.35 \pm 0.13 \pm 0.02$	AALTONEN	11L CDF	Repl. by AALTONEN 11AI

WEIGHTED AVERAGE  
 $0.52 \pm 0.16$  (Error scaled by 2.1)

 $B(B \rightarrow K\ell^+\ell^-)$  ( $1.0 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.33 \pm 0.13</math> OUR AVERAGE</b>			
$1.36^{+0.27}_{-0.24} \pm 0.03$	LEES	12s BABR	$e^+e^- \rightarrow \Upsilon(4S)$
$1.29 \pm 0.18 \pm 0.08$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
$1.36^{+0.23}_{-0.21} \pm 0.08$	WEI	09A BELL	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.01 \pm 0.26 \pm 0.07$	AALTONEN	11L CDF	Repl. by AALTONEN 11AI

 $B(B \rightarrow K\ell^+\ell^-)$  ( $0.0 < q^2 < 4.3 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.07 \pm 0.17 \pm 0.07</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.96 \pm 0.25 \pm 0.06$	AALTONEN	11L CDF	Repl. by AALTONEN 11AI

 $B(B \rightarrow X_s \ell^+ \ell^-)$  ( $1.0 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.60^{+0.41+0.25}_{-0.39-0.22}</math></b>	<sup>1</sup> LEES	14D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

 $B(B \rightarrow X_s \ell^+ e^-)$  ( $1.0 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.93^{+0.47+0.28}_{-0.45-0.24}</math></b>	<sup>1</sup> LEES	14D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

 $B(B \rightarrow X_s \mu^+ \mu^-)$  ( $1.0 < q^2 < 6.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.66^{+0.82+0.31}_{-0.76-0.25}</math></b>	<sup>1</sup> LEES	14D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

 $B(B \rightarrow X_s \ell^+ \ell^-)$  ( $14.2 < q^2 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.57^{+0.16+0.03}_{-0.15-0.02}</math></b>	<sup>1</sup> LEES	14D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

 $B(B \rightarrow X_s \ell^+ e^-)$  ( $14.2 < q^2 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.56^{+0.19+0.03}_{-0.18-0.03}</math></b>	<sup>1</sup> LEES	14D BABR	$e^+e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+ + \pi^-$  corrected for unobserved modes.

See key on page 601

# Meson Particle Listings

## $B^\pm/B^0$ ADMIXTURE

### $B(B \rightarrow X_s \mu^+ \mu^-) (14.2 < q^2 \text{ GeV}^2/c^4)$

VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT
$0.60^{+0.31+0.05}_{-0.29-0.04}$	<sup>1</sup> LEES	14D BABR	$e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Measured from sum of exclusive modes through  $K^+$ ,  $K^+\pi^0$ ,  $K^+\pi^-$ ,  $K^+\pi^-\pi^0$ ,  $K^+\pi^-\pi^+$ ,  $K_S^0$ ,  $K_S^0\pi^0$ ,  $K_S^0\pi^+$ ,  $K_S^0\pi^+\pi^0$ , and  $K_S^0\pi^+\pi^-\pi^0$  corrected for unobserved modes.

### LEPTON (HADRON) FORWARD-BACKWARD ASYMMETRY IN $B \rightarrow K^{(*)} \ell^+ \ell^- (B \rightarrow K/\pi h^+ h^-)$ DECAY

The forward-backward angular asymmetry of the lepton pair in  $B \rightarrow K^{(*)} \ell^+ \ell^- (B \rightarrow K/\pi h^+ h^-)$  decay is defined as

$$A_{FB}(s) = \frac{N(\cos\theta > 0) - N(\cos\theta < 0)}{N(\cos\theta > 0) + N(\cos\theta < 0)},$$

where  $s=q^2/m_B^2$ , and  $\theta$  is the angle of the  $\ell^- (h^-)$  with respect to the flight direction of the  $B$  meson, measured in the dilepton (dihadron) rest frame. In addition, the fraction of longitudinal polarization  $F_L$  of the  $K^*$  and  $F_S$ , the relative contribution from scalar and pseudoscalar penguin amplitudes in  $B \rightarrow K \ell^+ \ell^-$ , can be measured from the angular distribution of its decay products.

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (q^2 > 0.1 \text{ GeV}^2/c^4)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$0.50 \pm 0.15 \pm 0.02$		<sup>1</sup> SHIKAWA	06 BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$> 0.55$	95	<sup>2</sup> AUBERT,B	06J BABR	$e^+e^- \rightarrow \gamma(4S)$

<sup>1</sup> Using an unbinned max. likelihood fits to the  $M_{bc}$  distribution in five  $q^2$  bins for  $\cos\theta > 0$  and  $\cos\theta < 0$ .

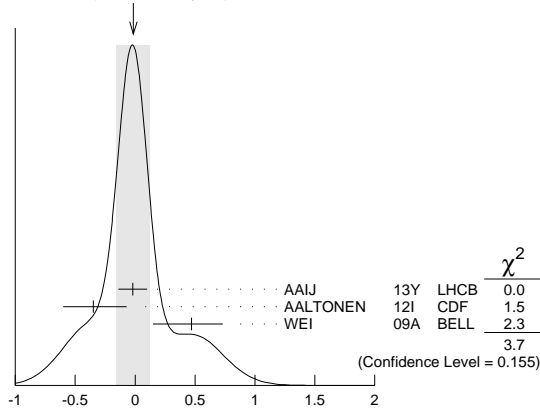
<sup>2</sup> Results with different  $q^2$  cuts are also reported.

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (0.1 < q^2 < 2.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.01 \pm 0.14$ OUR AVERAGE	Error includes scale factor of 1.4. See the ideogram below.		
$-0.02 \pm 0.12 \pm 0.01$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$-0.35^{+0.26}_{-0.23} \pm 0.10$	AALTONEN	12I CDF	$p\bar{p}$ at 1.96 TeV
$0.47^{+0.26}_{-0.32} \pm 0.03$	WEI	09A BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.29^{+0.37}_{-0.00} \pm 0.18$	<sup>1</sup> CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$-0.15 \pm 0.20 \pm 0.06$	AAIJ	12U LHCb	Repl. by AAIJ 13Y
$0.13^{+1.65}_{-0.75} \pm 0.25$	AALTONEN	11L CDF	Repl. by AALTONEN 12I

<sup>1</sup> CHATRCHYAN 13BL uses, for this bin,  $1.0 < q^2 < 2.0 \text{ GeV}^2/c^4$ .

WEIGHTED AVERAGE  
 $-0.01 \pm 0.14$  (Error scaled by 1.4)



$A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (0.1 < q^2 < 2.0 \text{ GeV}^2/c^4)$

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (m_{\ell\ell} < 2.5 \text{ GeV}^2/c^2)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.24^{+0.18}_{-0.23} \pm 0.05$	AUBERT	09N BABR	$e^+e^- \rightarrow \gamma(4S)$

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (m_{\ell\ell} > 3.2 \text{ GeV}^2/c^2)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.76^{+0.52}_{-0.32} \pm 0.07$	AUBERT	09N BABR	$e^+e^- \rightarrow \gamma(4S)$

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (0.10 < q^2 < 0.98 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.003^{+0.059}_{-0.057} \pm 0.009$	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (1.1 < q^2 < 2.5 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.191^{+0.068}_{-0.080} \pm 0.012$	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.14 \pm 0.05$ OUR AVERAGE			
$-0.118^{+0.082}_{-0.090} \pm 0.007$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$-0.12^{+0.15}_{-0.17} \pm 0.05$	KHACHATRY..16D	CMS	$pp$ at 8 TeV
$-0.20 \pm 0.08 \pm 0.01$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$-0.07 \pm 0.20 \pm 0.02$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.29^{+0.32}_{-0.35} \pm 0.15$	AALTONEN	12I CDF	$p\bar{p}$ at 1.96 TeV
$0.11^{+0.31}_{-0.36} \pm 0.07$	WEI	09A BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.05^{+0.16}_{-0.20} \pm 0.04$	AAIJ	12U LHCb	Repl. by AAIJ 13Y
$0.19^{+0.40}_{-0.41} \pm 0.14$	AALTONEN	11L CDF	Repl. by AALTONEN 12I

<sup>1</sup> Measured in  $2.5 < q^2 < 4.0 \text{ GeV}^2/c^4$ .

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08^{+0.21}_{-0.20} \pm 0.05$	AALTONEN	12I CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.21^{+0.31}_{-0.33} \pm 0.05$	AALTONEN	11L CDF	Repl. by AALTONEN 12I

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (4.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.025^{+0.051}_{-0.052} \pm 0.004$	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (6.0 < q^2 < 8.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.152^{+0.041}_{-0.040} \pm 0.008$	AAIJ	16B LHCb	$pp$ at 7, 8 TeV

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.092 \pm 0.027$ OUR AVERAGE			
$-0.075^{+0.032}_{-0.034} \pm 0.007$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$-0.12 \pm 0.08$	KHACHATRY..16D	CMS	$pp$ at 7, 8 TeV
$-0.17 \pm 0.06 \pm 0.01$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.29^{+0.20}_{-0.23} \pm 0.07$	AALTONEN	12I CDF	$p\bar{p}$ at 1.96 TeV
$0.26^{+0.27}_{-0.30} \pm 0.07$	WEI	09A BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.07 \pm 0.12 \pm 0.01$	CHATRCHYAN13BL	CMS	Repl. by KHACHATRYAN 16D
$-0.06^{+0.13}_{-0.14} \pm 0.07$	AAIJ	12U LHCb	Repl. by AAIJ 13Y
$0.43^{+0.36}_{-0.37} \pm 0.06$	AALTONEN	11L CDF	Repl. by AALTONEN 12I

<sup>1</sup> Measured in  $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ .

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (4.3 < q^2 < 8.6 \text{ GeV}^2/c^4)$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.13^{+0.06}_{-0.05} \pm 0.05$ OUR AVERAGE	Error includes scale factor of 1.1.		
$0.16^{+0.06}_{-0.05} \pm 0.01$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$-0.01 \pm 0.11 \pm 0.03$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.01 \pm 0.20 \pm 0.09$	AALTONEN	12I CDF	$p\bar{p}$ at 1.96 TeV
$0.45^{+0.15}_{-0.21} \pm 0.15$	WEI	09A BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.27^{+0.06}_{-0.08} \pm 0.02$	AAIJ	12U LHCb	Repl. by AAIJ 13Y
$-0.06^{+0.30}_{-0.28} \pm 0.05$	AALTONEN	11L CDF	Repl. by AALTONEN 12I

### $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$

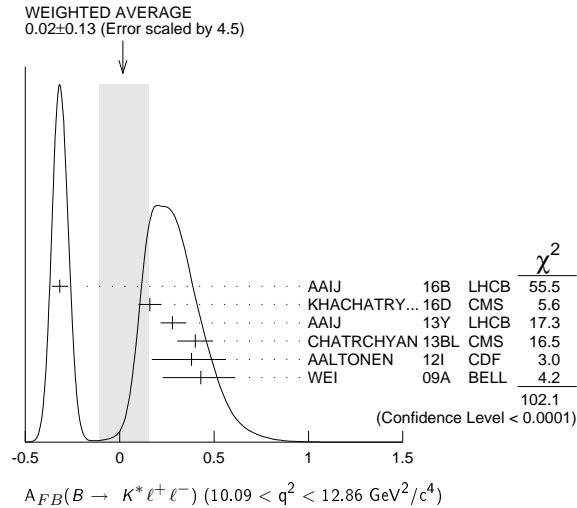
VALUE	DOCUMENT ID	TECN	COMMENT
$0.02 \pm 0.13$ OUR AVERAGE	Error includes scale factor of 4.5. See the ideogram below.		
$-0.318^{+0.044}_{-0.040} \pm 0.009$	<sup>1</sup> AAIJ	16B LHCb	$pp$ at 7, 8 TeV
$0.16 \pm 0.06 \pm 0.01$	KHACHATRY..16D	CMS	$pp$ at 8 TeV
$0.28^{+0.07}_{-0.06} \pm 0.02$	AAIJ	13Y LHCb	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.40 \pm 0.08 \pm 0.05$	CHATRCHYAN13BL	CMS	$pp$ at 7 TeV
$0.38^{+0.16}_{-0.19} \pm 0.09$	AALTONEN	12I CDF	$p\bar{p}$ at 1.96 TeV
$0.43^{+0.18}_{-0.20} \pm 0.03$	WEI	09A BELL	$e^+e^- \rightarrow \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

## Meson Particle Listings

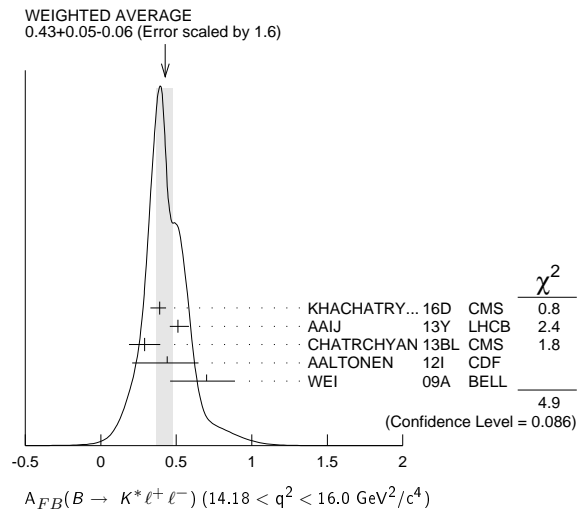
 $B^\pm/B^0$  ADMIXTURE

$0.27^{+0.11}_{-0.13} \pm 0.02$	AAIJ	12U	LHCB	Repl. by AAIJ 13Y
$0.66^{+0.23}_{-0.20} \pm 0.07$	AALTONEN	11L	CDF	Repl. by AALTONEN 12i

<sup>1</sup> Measured in  $11.0 < q^2 < 12.5 \text{ GeV}^2/c^4$ .

 $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (14.18 < q^2 < 16.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.43^{+0.05}_{-0.06} \pm 0.01</math> OUR AVERAGE</b>	Error includes scale factor of 1.6. See the ideogram below.		
$0.39^{+0.04}_{-0.06} \pm 0.01$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$0.51^{+0.07}_{-0.05} \pm 0.02$	AAIJ	13Y LHCB	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.29 \pm 0.09 \pm 0.05$	CHATRCHYAN 13BL	CMS	$pp$ at 7 TeV
$0.44^{+0.18}_{-0.21} \pm 0.10$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.70^{+0.16}_{-0.22} \pm 0.10$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.47^{+0.06}_{-0.08} \pm 0.03$	AAIJ	12U LHCB	Repl. by AAIJ 13Y
$0.42 \pm 0.16 \pm 0.09$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

 $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (15.0 < q^2 < 17.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.411^{+0.41}_{-0.037} \pm 0.008</math></b>	AAIJ	16B LHCB	$pp$ at 7, 8 TeV

 $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (17.0 < q^2 < 19.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.305^{+0.049}_{-0.048} \pm 0.013</math></b>	AAIJ	16B LHCB	$pp$ at 7, 8 TeV

 $A_{FB}(B \rightarrow K^* \ell^+ \ell^-) (16.0 < q^2 < 19.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.367 \pm 0.024</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.		
$0.355 \pm 0.027 \pm 0.009$	<sup>1</sup> AAIJ	16B LHCB	$pp$ at 7, 8 TeV
$0.35 \pm 0.07 \pm 0.01$	KHACHATRY...16D	CMS	$pp$ at 8 TeV
$0.30 \pm 0.08^{+0.01}_{-0.02}$	AAIJ	13Y LHCB	$pp$ at 7 TeV, $K^{*0} \mu^+ \mu^-$
$0.41 \pm 0.05 \pm 0.03$	CHATRCHYAN 13BL	CMS	$pp$ at 7 TeV
$0.65^{+0.17}_{-0.18} \pm 0.16$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.66^{+0.11}_{-0.16} \pm 0.04$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.16^{+0.11}_{-0.13} \pm 0.06$	AAIJ	12U LHCB	Repl. by AAIJ 13Y
$0.70^{+0.16}_{-0.25} \pm 0.10$	AALTONEN	11L CDF	Repl. by AALTONEN 12i
<sup>1</sup> Measured in $15.0 < q^2 < 19.0 \text{ GeV}^2/c^4$ .			

 $A_{FB}(B \rightarrow K \ell^+ \ell^-) (q^2 > 0.1 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.11 \pm 0.12</math> OUR AVERAGE</b>			
$0.15^{+0.21}_{-0.23} \pm 0.08$	<sup>1</sup> AUBERT, B	06J BABR	$e^+ e^- \rightarrow \Upsilon(4S)$
$0.10 \pm 0.14 \pm 0.01$	<sup>2</sup> ISHIKAWA	06 BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
<sup>1</sup> Results with different $q^2$ cuts are also reported.			
<sup>2</sup> Using an unbinned max. likelihood fits to the $M_{bc}$ distribution in five $q^2$ bins for $\cos \theta > 0$ and $\cos \theta < 0$ .			

 $A_{FB}(B \rightarrow K \ell^+ \ell^-) (q^2 < 2.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00^{+0.06}_{-0.05} \pm 0.03</math> OUR AVERAGE</b>			
$0.00^{+0.06}_{-0.05} \pm 0.03$	AAIJ	13H LHCB	$pp$ at 7 TeV
$0.13^{+0.42}_{-0.43} \pm 0.07$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$0.06^{+0.32}_{-0.35} \pm 0.02$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.15^{+0.46}_{-0.39} \pm 0.08$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

 $A_{FB}(B \rightarrow K \ell^+ \ell^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.09^{+0.10}_{-0.07} \pm 0.07</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.		
$0.07^{+0.08}_{-0.05} \pm 0.02$	AAIJ	13H LHCB	$pp$ at 7 TeV
$0.32^{+0.15}_{-0.16} \pm 0.05$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$-0.43^{+0.38}_{-0.40} \pm 0.09$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.72^{+0.40}_{-0.35} \pm 0.07$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

 $A_{FB}(B \rightarrow K \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.31 \pm 0.16 \pm 0.04</math></b>	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.36^{+0.24}_{-0.26} \pm 0.06$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

 $A_{FB}(B \rightarrow K \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.034^{+0.040}_{-0.029} \pm 0.029</math> OUR AVERAGE</b>			
$0.02^{+0.05}_{-0.03} \pm 0.02$	AAIJ	13H LHCB	$pp$ at 7 TeV
$0.13 \pm 0.09 \pm 0.02$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$-0.04^{+0.13}_{-0.16} \pm 0.05$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.08^{+0.27}_{-0.22} \pm 0.07$	AALTONEN	11L CDF	Repl. by AALTONEN 12i

 $A_{FB}(B \rightarrow K \ell^+ \ell^-) (4.3 < q^2 < 8.6 \text{ GeV}^2/c^4)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.04^{+0.04}_{-0.05} \pm 0.05</math> OUR AVERAGE</b>			
$-0.02^{+0.03}_{-0.05} \pm 0.03$	AAIJ	13H LHCB	$pp$ at 7 TeV
$0.01^{+0.13}_{-0.10} \pm 0.01$	AALTONEN	12i CDF	$p\bar{p}$ at 1.96 TeV
$-0.20^{+0.12}_{-0.14} \pm 0.03$	WEI	09A BELL	$e^+ e^- \rightarrow \Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.20^{+0.17}_{-0.28} \pm 0.03$	AALTONEN	11L CDF	Repl. by AALTONEN 12i









# Meson Particle Listings

## $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE

0.98 ±0.12 ±0.13	ONG	89	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
1.17 $\begin{smallmatrix} +0.27 \\ -0.22 \end{smallmatrix}$ $\begin{smallmatrix} +0.17 \\ -0.16 \end{smallmatrix}$	KLEM	88	DLCO	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
1.29 ±0.20 ±0.21	25 ASH	87	MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
1.02 $\begin{smallmatrix} +0.42 \\ -0.39 \end{smallmatrix}$	301 26 BROM	87	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
1 Measurement performed using an inclusive reconstruction and $B$ flavor identification technique.				
2 Measured using inclusive $J/\psi(1S) \rightarrow \mu^+ \mu^-$ vertex.				
3 ACCIARRI 98 uses inclusively reconstructed secondary vertex and lepton impact parameter.				
4 ACKERSTAFF 97F uses inclusively reconstructed secondary vertices.				
5 Combines ABREU 96E secondary vertex result with ABREU 94L impact parameter result.				
6 BUSKULIC 96F analyzed using 3D impact parameter.				
7 ABE,K 95B uses an inclusive topological technique.				
8 ABREU 94L uses charged particle impact parameters. Their result from inclusively reconstructed secondary vertices is superseded by ABREU 96E.				
9 From proper time distribution of $b \rightarrow J/\psi(1S)$ anything.				
10 ABE 93J analyzed using $J/\psi(1S) \rightarrow \mu \mu$ vertices.				
11 ABREU 93D data analyzed using $D/D^* \ell$ anything event vertices.				
12 ABREU 93G data analyzed using charged and neutral vertices.				
13 ACTON 93C analysed using $D/D^* \ell$ anything event vertices.				
14 ACTON 93L and ADRIANI 93K analyzed using lepton ( $e$ and $\mu$ ) impact parameter at Z.				
15 BUSKULIC 93O analyzed using dipole method.				
16 ABREU 92 is combined result of muon and hadron impact parameter analyses. Hadron tracks gave $(12.7 \pm 0.4 \pm 1.2) \times 10^{-13}$ s for an admixture of $B$ species weighted by production fraction and mean charge multiplicity, while muon tracks gave $(13.0 \pm 1.0 \pm 0.8) \times 10^{-13}$ s for an admixture weighted by production fraction and semileptonic branching fraction.				
17 ACTON 92 is combined result of muon and electron impact parameter analyses.				
18 BUSKULIC 92F uses the lepton impact parameter distribution for data from the 1991 run.				
19 BUSKULIC 92G use $J/\psi(1S)$ tags to measure the average $b$ lifetime. This is comparable to other methods only if the $J/\psi(1S)$ branching fractions of the different $b$ -flavored hadrons are in the same ratio.				
20 Using $Z \rightarrow e^+ X$ or $\mu^+ X$ , ADEVA 91H determined the average lifetime for an admixture of $B$ hadrons from the impact parameter distribution of the lepton.				
21 Using $Z \rightarrow J/\psi(1S) X$ , $J/\psi(1S) \rightarrow \ell^+ \ell^-$ , ALEXANDER 91G determined the average lifetime for an admixture of $B$ hadrons from the decay point of the $J/\psi(1S)$ .				
22 Using $Z \rightarrow e X$ or $\mu X$ , DECAMP 91C determines the average lifetime for an admixture of $B$ hadrons from the signed impact parameter distribution of the lepton.				
23 HAGEMANN 90 uses electrons and muons in an impact parameter analysis.				
24 LYONS 90 combine the results of the $B$ lifetime measurements of ONG 89, BRAUN-SCHWEIG 89B, KLEM 88, and ASH 87, and JADE data by private communication. They use statistical techniques which include variation of the error with the mean life, and possible correlations between the systematic errors. This result is not independent of the measured results used in our average.				
25 We have combined an overall scale error of 15% in quadrature with the systematic error of $\pm 0.7$ to obtain $\pm 2.1$ systematic error.				
26 Statistical and systematic errors were combined by BROM 87.				

### CHARGED $b$ -HADRON ADMIXTURE MEAN LIFE

VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
<b>1.72±0.08±0.06</b>	1 ADAM	95	DLPH $e^+ e^- \rightarrow Z$
1 ADAM 95 data analyzed using vertex-charge technique to tag $b$ -hadron charge.			

### NEUTRAL $b$ -HADRON ADMIXTURE MEAN LIFE

VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
<b>1.58±0.11±0.09</b>	1 ADAM	95	DLPH $e^+ e^- \rightarrow Z$
1 ADAM 95 data analyzed using vertex-charge technique to tag $b$ -hadron charge.			

### MEAN LIFE RATIO $\tau_{\text{charged } b\text{-hadron}}/\tau_{\text{neutral } b\text{-hadron}}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.09<math>\begin{smallmatrix} +0.11 \\ -0.10 \end{smallmatrix}</math>±0.08</b>	1 ADAM	95	DLPH $e^+ e^- \rightarrow Z$
1 ADAM 95 data analyzed using vertex-charge technique to tag $b$ -hadron charge.			

$$|\Delta\tau_b|/\tau_{b,\overline{b}}$$

$\tau_{b,\overline{b}}$  and  $|\Delta\tau_b|$  are the mean life average and difference between  $b$  and  $\overline{b}$  hadrons.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.001±0.012±0.008</b>	1 ABBIENDI	99J	OPAL $e^+ e^- \rightarrow Z$
1 Data analyzed using both the jet charge and the charge of secondary vertex in the opposite hemisphere.			

### $\overline{b}$ PRODUCTION FRACTIONS AND DECAY MODES

The branching fraction measurements are for an admixture of  $B$  mesons and baryons at energies above the  $T(4S)$ . Only the highest energy results (LHC, LEP, Tevatron,  $Sp\overline{p}S$ ) are used in the branching fraction averages. In the following, we assume that the production fractions are the same at the LHC, LEP, and at the Tevatron.

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm$  anything, the values usually are multiplicities, not branching fractions. They can be greater than one.

The modes below are listed for a  $\overline{b}$  initial state.  $b$  modes are their charge conjugates. Reactions indicate the weak decay vertex and do not include mixing.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
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### PRODUCTION FRACTIONS

The production fractions for weakly decaying  $b$ -hadrons at high energy have been calculated from the best values of mean lives, mixing parameters, and branching fractions in this edition by the Heavy Flavor Averaging Group (HFAG) as described in the note “ $B^0$ - $\overline{B}^0$  Mixing” in the  $B^0$  Particle Listings. The production fractions in  $b$ -hadronic  $Z$  decay or  $p\overline{p}$  collisions at the Tevatron are also listed at the end of the section. Values assume

$$\begin{aligned} B(\overline{b} \rightarrow B^+) &= B(\overline{b} \rightarrow B^0) \\ B(\overline{b} \rightarrow B^+) + B(\overline{b} \rightarrow B^0) + B(\overline{b} \rightarrow B_s^0) + B(b \rightarrow b\text{-baryon}) &= 100\%. \end{aligned}$$

The correlation coefficients between production fractions are also reported:

$$\begin{aligned} \text{cor}(B_s^0, b\text{-baryon}) &= -0.240 \\ \text{cor}(B_s^0, B^\pm=B^0) &= -0.161 \\ \text{cor}(b\text{-baryon}, B^\pm=B^0) &= -0.920. \end{aligned}$$

The notation for production fractions varies in the literature ( $f_d$ ,  $d_{B^0}$ ,  $f(b \rightarrow \overline{B}^0)$ ,  $\text{Br}(b \rightarrow \overline{B}^0)$ ). We use our own branching fraction notation here,  $B(\overline{b} \rightarrow B^0)$ .

Note these production fractions are  $b$ -hadronization fractions, not the conventional branching fractions of  $b$ -quark to a  $B$ -hadron, which may have considerable dependence on the initial and final state kinematic and production environment.

$\Gamma_1$	$B^+$	( 40.4 ± 0.6 ) %
$\Gamma_2$	$B^0$	( 40.4 ± 0.6 ) %
$\Gamma_3$	$B_s^0$	( 10.3 ± 0.5 ) %
$\Gamma_4$	$b$ -baryon	( 8.9 ± 1.3 ) %

### DECAY MODES

#### Semileptonic and leptonic modes

$\Gamma_5$	$\nu$ anything	( 23.1 ± 1.5 ) %	
$\Gamma_6$	$\ell^+ \nu_\ell$ anything	[a] ( 10.69± 0.22 ) %	
$\Gamma_7$	$e^+ \nu_e$ anything	( 10.86± 0.35 ) %	
$\Gamma_8$	$\mu^+ \nu_\mu$ anything	( 10.95 $\begin{smallmatrix} + \\ - \end{smallmatrix}$ 0.29 $\begin{smallmatrix} + \\ - \end{smallmatrix}$ 0.25 ) %	
$\Gamma_9$	$D^- \ell^+ \nu_\ell$ anything	[a] ( 2.2 ± 0.4 ) %	S=1.9
$\Gamma_{10}$	$D^- \pi^+ \ell^+ \nu_\ell$ anything	( 4.9 ± 1.9 ) $\times 10^{-3}$	
$\Gamma_{11}$	$D^- \pi^- \ell^+ \nu_\ell$ anything	( 2.6 ± 1.6 ) $\times 10^{-3}$	
$\Gamma_{12}$	$\overline{D}^0 \ell^+ \nu_\ell$ anything	[a] ( 6.81± 0.34 ) %	
$\Gamma_{13}$	$\overline{D}^0 \pi^- \ell^+ \nu_\ell$ anything	( 1.07± 0.27 ) %	
$\Gamma_{14}$	$\overline{D}^0 \pi^+ \ell^+ \nu_\ell$ anything	( 2.3 ± 1.6 ) $\times 10^{-3}$	
$\Gamma_{15}$	$D^{*-} \ell^+ \nu_\ell$ anything	[a] ( 2.75± 0.19 ) %	
$\Gamma_{16}$	$D^{*-} \pi^- \ell^+ \nu_\ell$ anything	( 6 ± 7 ) $\times 10^{-4}$	
$\Gamma_{17}$	$D^{*-} \pi^+ \ell^+ \nu_\ell$ anything	( 4.8 ± 1.0 ) $\times 10^{-3}$	
$\Gamma_{18}$	$\overline{D}_j^0 \ell^+ \nu_\ell$ anything $\times$ $B(\overline{D}_j^0 \rightarrow D^{*+} \pi^-)$	[a,b] ( 2.6 ± 0.9 ) $\times 10^{-3}$	
$\Gamma_{19}$	$D_j^- \ell^+ \nu_\ell$ anything $\times$ $B(D_j^- \rightarrow D^0 \pi^-)$	[a,b] ( 7.0 ± 2.3 ) $\times 10^{-3}$	
$\Gamma_{20}$	$\overline{D}_2^*(2460)^0 \ell^+ \nu_\ell$ anything $\times B(\overline{D}_2^*(2460)^0 \rightarrow$ $D^{*-} \pi^+)$	< 1.4 $\times 10^{-3}$	CL=90%
$\Gamma_{21}$	$D_2^*(2460)^- \ell^+ \nu_\ell$ anything $\times B(D_2^*(2460)^- \rightarrow$ $D^0 \pi^-)$	( 4.2 $\begin{smallmatrix} + \\ - \end{smallmatrix}$ 1.5 $\begin{smallmatrix} + \\ - \end{smallmatrix}$ 1.8 ) $\times 10^{-3}$	
$\Gamma_{22}$	$\overline{D}_2^*(2460)^0 \ell^+ \nu_\ell$ anything $\times B(\overline{D}_2^*(2460)^0 \rightarrow$ $D^- \pi^+)$	( 1.6 ± 0.8 ) $\times 10^{-3}$	

See key on page 601

# Meson Particle Listings

## $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE

$\Gamma_{23}$	charmless $\ell\bar{\nu}_\ell$	[a]	$(1.7 \pm 0.5) \times 10^{-3}$
$\Gamma_{24}$	$\tau^+ \nu_\tau$ anything		$(2.41 \pm 0.23) \%$
$\Gamma_{25}$	$D^{*-} \tau \nu_\tau$ anything		$(9 \pm 4) \times 10^{-3}$
$\Gamma_{26}$	$\bar{c} \rightarrow \ell^- \bar{\nu}_\ell$ anything	[a]	$(8.02 \pm 0.19) \%$
$\Gamma_{27}$	$c \rightarrow \ell^+ \nu$ anything		$(1.6 \pm_{-0.5}^{+0.4}) \%$

### Charmed meson and baryon modes

$\Gamma_{28}$	$\bar{D}^0$ anything		$(59.0 \pm 2.9) \%$
$\Gamma_{29}$	$D^0 D_s^\pm$ anything	[c]	$(9.1 \pm_{-2.8}^{+4.0}) \%$
$\Gamma_{30}$	$D^\mp D_s^\pm$ anything	[c]	$(4.0 \pm_{-1.8}^{+2.3}) \%$
$\Gamma_{31}$	$\bar{D}^0 D^0$ anything	[c]	$(5.1 \pm_{-1.8}^{+2.0}) \%$
$\Gamma_{32}$	$D^0 D^\pm$ anything	[c]	$(2.7 \pm_{-1.6}^{+1.8}) \%$
$\Gamma_{33}$	$D^\pm D^\mp$ anything	[c]	$< 9 \times 10^{-3}$ CL=90%
$\Gamma_{34}$	$D^0$ anything		
$\Gamma_{35}$	$D^+$ anything		
$\Gamma_{36}$	$D^-$ anything		$(22.5 \pm 1.7) \%$
$\Gamma_{37}$	$D^*(2010)^+$ anything		$(17.3 \pm 2.0) \%$
$\Gamma_{38}$	$D_1(2420)^0$ anything		$(5.0 \pm 1.5) \%$
$\Gamma_{39}$	$D^*(2010)^\mp D_s^\pm$ anything	[c]	$(3.3 \pm_{-1.3}^{+1.6}) \%$
$\Gamma_{40}$	$D^0 D^*(2010)^\pm$ anything	[c]	$(3.0 \pm_{-1.1}^{+0.9}) \%$
$\Gamma_{41}$	$D^*(2010)^\pm D^\mp$ anything	[c]	$(2.5 \pm_{-1.0}^{+1.2}) \%$
$\Gamma_{42}$	$D^*(2010)^\pm D^*(2010)^\mp$ anything	[c]	$(1.2 \pm 0.4) \%$
$\Gamma_{43}$	$\bar{D} D$ anything		$(10 \pm_{-10}^{+11}) \%$
$\Gamma_{44}$	$D_s^*(2460)^0$ anything		$(4.7 \pm 2.7) \%$
$\Gamma_{45}$	$D_s^-$ anything		$(14.7 \pm 2.1) \%$
$\Gamma_{46}$	$D_s^+$ anything		$(10.1 \pm 3.1) \%$
$\Gamma_{47}$	$\Lambda_c^+$ anything		$(7.6 \pm 1.1) \%$
$\Gamma_{48}$	$\bar{c}/c$ anything	[d]	$(116.2 \pm 3.2) \%$

### Charmonium modes

$\Gamma_{49}$	$J/\psi(1S)$ anything		$(1.16 \pm 0.10) \%$
$\Gamma_{50}$	$\psi(2S)$ anything		$(2.83 \pm 0.29) \times 10^{-3}$
$\Gamma_{51}$	$\chi_{c1}(1P)$ anything		$(1.4 \pm 0.4) \%$

### K or $K^*$ modes

$\Gamma_{52}$	$\bar{S}\gamma$		$(3.1 \pm 1.1) \times 10^{-4}$
$\Gamma_{53}$	$\bar{S}\bar{\nu}\nu$	B1	$< 6.4 \times 10^{-4}$ CL=90%
$\Gamma_{54}$	$K^\pm$ anything		$(74 \pm 6) \%$
$\Gamma_{55}$	$K_S^0$ anything		$(29.0 \pm 2.9) \%$

### Pion modes

$\Gamma_{56}$	$\pi^\pm$ anything		$(397 \pm 21) \%$
$\Gamma_{57}$	$\pi^0$ anything	[d]	$(278 \pm 60) \%$
$\Gamma_{58}$	$\phi$ anything		$(2.82 \pm 0.23) \%$

### Baryon modes

$\Gamma_{59}$	$p/\bar{p}$ anything		$(13.1 \pm 1.1) \%$
$\Gamma_{60}$	$\Lambda/\bar{\Lambda}$ anything		$(5.9 \pm 0.6) \%$
$\Gamma_{61}$	$b$ -baryon anything		$(10.2 \pm 2.8) \%$

### Other modes

$\Gamma_{62}$	charged anything	[d]	$(497 \pm 7) \%$
$\Gamma_{63}$	hadron <sup>+</sup> hadron <sup>-</sup>		$(1.7 \pm_{-0.7}^{+1.0}) \times 10^{-5}$
$\Gamma_{64}$	charmless		$(7 \pm 21) \times 10^{-3}$

### $\Delta B = 1$ weak neutral current (B1) modes

$\Gamma_{65}$	$e^+ e^-$ anything	B1	
$\Gamma_{66}$	$\mu^+ \mu^-$ anything	B1	$< 3.2 \times 10^{-4}$ CL=90%
$\Gamma_{67}$	$\nu\bar{\nu}$ anything	B1	

[a] An  $\ell$  indicates an  $e$  or a  $\mu$  mode, not a sum over these modes.[b]  $D_j$  represents an unresolved mixture of pseudoscalar and tensor  $D^{**}$  ( $P$ -wave) states.

[c] The value is for the sum of the charge states or particle/antiparticle states indicated.

[d] Inclusive branching fractions have a multiplicity definition and can be greater than 100%.

### $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE BRANCHING RATIOS

$\Gamma(B^+)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
"OUR EVALUATION" is an average using rescaled values of the data listed below and from the best values of mean lives, mixing parameters, and branching fractions in this edition by the Heavy Flavor Averaging Group (HFAG) as described at <a href="http://www.slac.stanford.edu/xorg/hfag/">http://www.slac.stanford.edu/xorg/hfag/</a> .	

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.405 <math>\pm 0.006</math> OUR EVALUATION</b>			
<b>0.4099 <math>\pm 0.0082 \pm 0.0111</math></b>	<sup>1</sup> ABDALLAH	03k	DLPB $e^+e^- \rightarrow Z$

<sup>1</sup> The analysis is based on a neural network, to estimate the charge of the weakly-decaying  $b$  hadron by distinguishing its decay products from particles produced at the primary vertex.

$\Gamma(B^+)/\Gamma(B^0)$	$\Gamma_1/\Gamma_2$		
VALUE	DOCUMENT ID	TECN	COMMENT
$1.054 \pm 0.018^{+0.062}_{-0.074}$	AALTONEN	08N CDF	$p\bar{p}$ at 1.96 TeV

$\Gamma(B_s^0)/[\Gamma(B^+) + \Gamma(B^0)]$	$\Gamma_3/(\Gamma_1 + \Gamma_2)$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.130 <math>\pm 0.008</math> OUR EVALUATION</b>			
<b>0.134 <math>\pm 0.008</math> OUR AVERAGE</b>			

0.134 $\pm 0.004 \pm_{-0.010}^{+0.011}$	<sup>1</sup> AAIJ	12J	LHCB $pp$ at 7 TeV
0.1265 $\pm 0.0085 \pm 0.0131$	<sup>2</sup> AAIJ	11F	LHCB $pp$ at 7 TeV
0.128 $\pm_{-0.010}^{+0.011}$	<sup>3</sup> AALTONEN	08N	CDF $p\bar{p}$ at 1.96 TeV
0.213 $\pm 0.068$	<sup>4</sup> AFFOLDER	00E	CDF $p\bar{p}$ at 1.8 TeV
0.21 $\pm_{-0.030}^{+0.036}$	<sup>5</sup> ABE	99P	CDF $p\bar{p}$ at 1.8 TeV

<sup>1</sup> Measured using  $b$ -hadron semileptonic decays and assuming isospin symmetry.

<sup>2</sup> AAIJ 11F measured  $f_s/f_d = 0.253 \pm 0.017 \pm 0.017 \pm 0.020$ , where the errors are statistical, systematic, and theoretical. We divide their value by 2. Our second error combines systematic and theoretical uncertainties.

<sup>3</sup> AALTONEN 08N reports  $[\Gamma(\bar{b} \rightarrow B_s^0)/[\Gamma(\bar{b} \rightarrow B^+) + \Gamma(\bar{b} \rightarrow B^0)]] \times [B(D_s^+ \rightarrow \phi\pi^+)] = (5.76 \pm 0.18 \pm_{-0.42}^{+0.45}) \times 10^{-3}$  which we divide by our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> AFFOLDER 00E uses several electron-charm final states in  $b \rightarrow ce^-X$ .

<sup>5</sup> ABE 99P uses the numbers of  $K^*(892)^0$ ,  $K^*(892)^+$ , and  $\phi(1020)$  events produced in association with the double semileptonic decays  $b \rightarrow c\mu^-X$  with  $c \rightarrow s\mu^+X$ .

$\Gamma(B_s^0)/\Gamma(B^0)$	$\Gamma_3/\Gamma_2$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.256±0.014 OUR EVALUATION</b>			
<b>0.239±0.016 OUR AVERAGE</b>			

0.240 $\pm 0.004 \pm 0.020$	<sup>1</sup> AAD	15CMATLS	$pp$ at 7 TeV
0.238 $\pm 0.004 \pm 0.015 \pm 0.021$	<sup>2</sup> AAIJ	13P	LHCB $pp$ at 7 TeV

<sup>1</sup> The measurement is derived from the observed  $B_s^0 \rightarrow J/\psi\phi$  and  $B_d^0 \rightarrow J/\psi K^{*0}$  yields and a recent theory prediction of  $B(B_s^0 \rightarrow J/\psi\phi)/B(B_d^0 \rightarrow J/\psi K^{*0})$ . The second uncertainty combines in quadrature systematic and theoretical uncertainties.

<sup>2</sup> AAIJ 13P studies also separately the  $p_T(B)$  and  $\eta(B)$  dependency of  $\Gamma(\bar{b} \rightarrow B_s^0)/\Gamma(\bar{b} \rightarrow B^0)$ , finding  $f_s/f_d(p_T) = (0.256 \pm 0.020) + (-2.0 \pm 0.6) 10^{-3} / \text{GeV}/c$  ( $p_T - \langle p_T \rangle$ ) and  $f_s/f_d(\eta) = (0.256 \pm 0.020) + (0.005 \pm 0.006) (\eta - \langle \eta \rangle)$ , where  $\langle p_T \rangle = 10.4 \text{ GeV}/c$  and  $\langle \eta \rangle = 3.28$ .

$\Gamma(b\text{-baryon})/[\Gamma(B^+) + \Gamma(B^0)]$	$\Gamma_4/(\Gamma_1+\Gamma_2)$		
VALUE	DOCUMENT ID	TECN	COMMENT
0.105 ± 0.015 OUR EVALUATION			
0.27 $\begin{smallmatrix} +0.06 \\ -0.05 \end{smallmatrix}$ OUR AVERAGE			

0.305 $\pm 0.010 \pm 0.081$	<sup>1</sup> AAIJ	12J	LHCB $pp$ at 7 TeV
0.31 $\pm_{-0.08}^{+0.11}$	<sup>2</sup> AALTONEN	09E	CDF $p\bar{p}$ at 1.8 TeV
0.22 $\pm_{-0.07}^{+0.08}$ $\pm 0.01$	<sup>3</sup> AALTONEN	08N	CDF $p\bar{p}$ at 1.96 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.118 $\pm 0.042$	<sup>4</sup> AFFOLDER	00E	CDF $p\bar{p}$ at 1.8 TeV
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<sup>1</sup> Measured the ratio to be  $(0.404 \pm 0.017 \pm 0.027 \pm 0.105) \times [1 - (0.031 \pm 0.004 \pm 0.003) \times p_T]$  using  $b$ -hadron semileptonic decays where the  $p_T$  is the momentum of charmed hadron-muon pair in  $\text{GeV}/c$ . We quote their weighted average value where the second error combines systematic and the error on  $B(\Lambda_c^+ \rightarrow pK^- \pi^+)$ .

<sup>2</sup> Errata to the measurement reported in AFFOLDER 00E using the  $p_T$  spectra from fully reconstructed  $B^0$  and  $\Lambda_b$  decays.

<sup>3</sup> AALTONEN 08N reports  $[\Gamma(\bar{b} \rightarrow b\text{-baryon})/[\Gamma(\bar{b} \rightarrow B^+) + \Gamma(\bar{b} \rightarrow B^0)]] \times [B(\Lambda_c^+ \rightarrow pK^- \pi^+)] = (14.1 \pm 0.6 \pm_{-4.4}^{+5.3}) \times 10^{-3}$  which we divide by our best value  $B(\Lambda_c^+ \rightarrow pK^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> AFFOLDER 00E uses several electron-charm final states in  $b \rightarrow ce^-X$ .

# Meson Particle Listings

## $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE

$\Gamma(\nu\text{anything})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.2308<math>\pm</math>0.0077<math>\pm</math>0.0124</b>	1,2 ACCIARRI	96c	L3	$e^+e^- \rightarrow Z$

- <sup>1</sup> ACCIARRI 96c assumes relative  $b$  semileptonic decay rates  $e:\mu:\tau$  of 1:1:0.25. Based on missing-energy spectrum.  
<sup>2</sup> Assumes Standard Model value for  $R_B$ .

$\Gamma(\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
"OUR EVALUATION" is an average of the data listed below, excluding all asymmetry measurements, performed by the LEP Electroweak Working Group as described in the "Note on the Z boson" in the Z Particle Listings.				

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.1069<math>\pm</math>0.0022 OUR EVALUATION</b>			
<b>0.1064<math>\pm</math>0.0016 OUR AVERAGE</b>			
0.1070 $\pm$ 0.0010 $\pm$ 0.0035	<sup>1</sup> HEISTER	02G	ALEP $e^+e^- \rightarrow Z$
0.1070 $\pm$ 0.0008 $^{+0.0037}_{-0.0049}$	<sup>2</sup> ABREU	01L	DLPH $e^+e^- \rightarrow Z$
0.1083 $\pm$ 0.0010 $^{+0.0028}_{-0.0024}$	<sup>3</sup> ABBIENDI	00E	OPAL $e^+e^- \rightarrow Z$
0.1016 $\pm$ 0.0013 $\pm$ 0.0030	<sup>4</sup> ACCIARRI	00	L3 $e^+e^- \rightarrow Z$
0.1085 $\pm$ 0.0012 $\pm$ 0.0047	<sup>5,6</sup> ACCIARRI	96c	L3 $e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1106 $\pm$ 0.0039 $\pm$ 0.0022	<sup>7</sup> ABREU	95D	DLPH $e^+e^- \rightarrow Z$
0.114 $\pm$ 0.003 $\pm$ 0.004	<sup>8</sup> BUSKULIC	94G	ALEP $e^+e^- \rightarrow Z$
0.100 $\pm$ 0.007 $\pm$ 0.007	<sup>9</sup> ABREU	93c	DLPH $e^+e^- \rightarrow Z$
0.105 $\pm$ 0.006 $\pm$ 0.005	<sup>10</sup> AKERS	93B	OPAL Repl. by ABBI- ENDI 00E

- <sup>1</sup> Uses the combination of lepton transverse momentum spectrum and the correlation between the charge of the lepton and opposite jet charge. The first error is statistic and the second error is the total systematic error including the modeling.  
<sup>2</sup> The experimental systematic and model uncertainties are combined in quadrature.  
<sup>3</sup> ABBIENDI 00E result is determined by comparing the distribution of several kinematic variables of leptonic events in a lifetime tagged  $Z \rightarrow b\bar{b}$  sample using artificial neural network techniques. The first error is statistic; the second error is the total systematic error.  
<sup>4</sup> ACCIARRI 00 result obtained from a combined fit of  $R_D = \Gamma(Z \rightarrow b\bar{b})/\Gamma(Z \rightarrow \text{hadrons})$  and  $B(b \rightarrow \ell\nu X)$ , using double-tagging method.  
<sup>5</sup> ACCIARRI 96c result obtained by a fit to the single lepton spectrum.  
<sup>6</sup> Assumes Standard Model value for  $R_B$ .  
<sup>7</sup> ABREU 95D give systematic errors  $\pm 0.0019$  (model) and 0.0012 ( $R_C$ ). We combine these in quadrature.  
<sup>8</sup> BUSKULIC 94G uses  $e$  and  $\mu$  events. This value is from a global fit to the lepton  $p_T$  (relative to jet) spectra which also determines the  $b$  and  $c$  production fractions, the fragmentation functions, and the forward-backward asymmetries. This branching ratio depends primarily on the ratio of dileptons to single leptons at high  $p_T$ , but the lower  $p_T$  portion of the lepton spectrum is included in the global fit to reduce the model dependence. The model dependence is  $\pm 0.0026$  and is included in the systematic error.  
<sup>9</sup> ABREU 93c event count includes  $ee$  events. Combining  $ee$ ,  $\mu\mu$ , and  $e\mu$  events, they obtain  $0.100 \pm 0.007 \pm 0.007$ .  
<sup>10</sup> AKERS 93B analysis performed using single and dilepton events.

$\Gamma(e^+\nu_e\text{anything})/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1086<math>\pm</math>0.0035 OUR AVERAGE</b>				

0.1078 $\pm$ 0.0008 $^{+0.0050}_{-0.0046}$	<sup>1</sup> ABBIENDI	00E	OPAL	$e^+e^- \rightarrow Z$
0.1089 $\pm$ 0.0020 $\pm$ 0.0051	<sup>2,3</sup> ACCIARRI	96c	L3	$e^+e^- \rightarrow Z$
0.107 $\pm$ 0.015 $\pm$ 0.007	<sup>4</sup> ABREU	93c	DLPH	$e^+e^- \rightarrow Z$
0.138 $\pm$ 0.032 $\pm$ 0.008	<sup>5</sup> ADEVA	91c	L3	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.086 $\pm$ 0.027 $\pm$ 0.008	<sup>6</sup> ABE	93E	VNS	$E_{\text{CM}}^{\text{ee}} = 58 \text{ GeV}$
0.109 $^{+0.014}_{-0.013}$ $\pm$ 0.0055	<sup>7</sup> AKERS	93B	OPAL	Repl. by ABBI- ENDI 00E
0.111 $\pm$ 0.028 $\pm$ 0.026	BEHREND	90D	CELL	$E_{\text{CM}}^{\text{ee}} = 43 \text{ GeV}$
0.150 $\pm$ 0.011 $\pm$ 0.022	BEHREND	90D	CELL	$E_{\text{CM}}^{\text{ee}} = 35 \text{ GeV}$
0.112 $\pm$ 0.009 $\pm$ 0.011	ONG	88	MRK2	$E_{\text{CM}}^{\text{ee}} = 29 \text{ GeV}$
0.149 $^{+0.022}_{-0.019}$	PAL	86	DLCO	$E_{\text{CM}}^{\text{ee}} = 29 \text{ GeV}$
0.110 $\pm$ 0.018 $\pm$ 0.010	AIHARA	85	TPC	$E_{\text{CM}}^{\text{ee}} = 29 \text{ GeV}$
0.111 $\pm$ 0.034 $\pm$ 0.040	ALTHOFF	84J	TASS	$E_{\text{CM}}^{\text{ee}} = 34.6 \text{ GeV}$
0.146 $\pm$ 0.028	KOOP	84	DLCO	Repl. by PAL 86
0.116 $\pm$ 0.021 $\pm$ 0.017	NELSON	83	MRK2	$E_{\text{CM}}^{\text{ee}} = 29 \text{ GeV}$

- <sup>1</sup> ABBIENDI 00E result is determined by comparing the distribution of several kinematic variables of leptonic events in a lifetime tagged  $Z \rightarrow b\bar{b}$  sample using artificial neural network techniques. The first error is statistic; the second error is the total systematic error.  
<sup>2</sup> ACCIARRI 96c result obtained by a fit to the single lepton spectrum.  
<sup>3</sup> Assumes Standard Model value for  $R_B$ .  
<sup>4</sup> ABREU 93c event count includes  $ee$  events. Combining  $ee$ ,  $\mu\mu$ , and  $e\mu$  events, they obtain  $0.100 \pm 0.007 \pm 0.007$ .  
<sup>5</sup> ADEVA 91c measure the average  $B(b \rightarrow eX)$  branching ratio using single and double tagged  $b$  enhanced  $Z$  events. Combining  $e$  and  $\mu$  results, they obtain  $0.113 \pm 0.010 \pm 0.006$ . Constraining the initial number of  $b$  quarks by the Standard Model prediction ( $378 \pm 3 \text{ MeV}$ ) for the decay of the  $Z$  into  $b\bar{b}$ , the electron result gives  $0.112 \pm 0.004 \pm 0.008$ . They obtain  $0.119 \pm 0.003 \pm 0.006$  when  $e$  and  $\mu$  results are combined. Used to measure the  $b\bar{b}$  width itself, this electron result gives  $370 \pm 12 \pm 24 \text{ MeV}$  and combined with the muon result gives  $385 \pm 7 \pm 22 \text{ MeV}$ .  
<sup>6</sup> ABE 93E experiment also measures forward-backward asymmetries and fragmentation functions for  $b$  and  $c$ .  
<sup>7</sup> AKERS 93B analysis performed using single and dilepton events.

$\Gamma(\mu^+\nu_\mu\text{anything})/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1095<math>\pm</math>0.0029 -0.0025 OUR AVERAGE</b>				

0.1096 $\pm$ 0.0008 $^{+0.0034}_{-0.0027}$	<sup>1</sup> ABBIENDI	00E	OPAL	$e^+e^- \rightarrow Z$
0.1082 $\pm$ 0.0015 $\pm$ 0.0059	<sup>2,3</sup> ACCIARRI	96c	L3	$e^+e^- \rightarrow Z$
0.110 $\pm$ 0.012 $\pm$ 0.007	<sup>4</sup> ABREU	93c	DLPH	$e^+e^- \rightarrow Z$
0.113 $\pm$ 0.012 $\pm$ 0.006	<sup>5</sup> ADEVA	91c	L3	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.122 $\pm$ 0.006 $\pm$ 0.007	<sup>3</sup> UENO	96	AMY	$e^+e^-$ at 57.9 GeV
0.101 $^{+0.010}_{-0.009}$ $\pm$ 0.0055	<sup>6</sup> AKERS	93B	OPAL	Repl. by ABBI- ENDI 00E
0.104 $\pm$ 0.023 $\pm$ 0.016	BEHREND	90D	CELL	$E_{\text{CM}}^{\text{ee}} = 43 \text{ GeV}$
0.148 $\pm$ 0.010 $\pm$ 0.016	BEHREND	90D	CELL	$E_{\text{CM}}^{\text{ee}} = 35 \text{ GeV}$
0.118 $\pm$ 0.012 $\pm$ 0.010	ONG	88	MRK2	$E_{\text{CM}}^{\text{ee}} = 29 \text{ GeV}$
0.117 $\pm$ 0.016 $\pm$ 0.015	BARTEL	87	JADE	$E_{\text{CM}}^{\text{ee}} = 34.6 \text{ GeV}$
0.114 $\pm$ 0.018 $\pm$ 0.025	BARTEL	85J	JADE	Repl. by BARTEL 87
0.117 $\pm$ 0.028 $\pm$ 0.010	ALTHOFF	84G	TASS	$E_{\text{CM}}^{\text{ee}} = 34.5 \text{ GeV}$
0.105 $\pm$ 0.015 $\pm$ 0.013	ADEVA	83B	MRKJ	$E_{\text{CM}}^{\text{ee}} = 33\text{--}38.5 \text{ GeV}$
0.155 $^{+0.054}_{-0.029}$	FERNANDEZ	83D	MAC	$E_{\text{CM}}^{\text{ee}} = 29 \text{ GeV}$

- <sup>1</sup> ABBIENDI 00E result is determined by comparing the distribution of several kinematic variables of leptonic events in a lifetime tagged  $Z \rightarrow b\bar{b}$  sample using artificial neural network techniques. The first error is statistic; the second error is the total systematic error.  
<sup>2</sup> ACCIARRI 96c result obtained by a fit to the single lepton spectrum.  
<sup>3</sup> Assumes Standard Model value for  $R_B$ .  
<sup>4</sup> ABREU 93c event count includes  $\mu\mu$  events. Combining  $ee$ ,  $\mu\mu$ , and  $e\mu$  events, they obtain  $0.100 \pm 0.007 \pm 0.007$ .  
<sup>5</sup> ADEVA 91c measure the average  $B(b \rightarrow eX)$  branching ratio using single and double tagged  $b$  enhanced  $Z$  events. Combining  $e$  and  $\mu$  results, they obtain  $0.113 \pm 0.010 \pm 0.006$ . Constraining the initial number of  $b$  quarks by the Standard Model prediction ( $378 \pm 3 \text{ MeV}$ ) for the decay of the  $Z$  into  $b\bar{b}$ , the muon result gives  $0.123 \pm 0.003 \pm 0.006$ . They obtain  $0.119 \pm 0.003 \pm 0.006$  when  $e$  and  $\mu$  results are combined. Used to measure the  $b\bar{b}$  width itself, this muon result gives  $394 \pm 9 \pm 22 \text{ MeV}$  and combined with the electron result gives  $385 \pm 7 \pm 22 \text{ MeV}$ .  
<sup>6</sup> AKERS 93B analysis performed using single and dilepton events.

$\Gamma(D^-\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.022 <math>\pm</math>0.004 OUR AVERAGE</b>				Error includes scale factor of 1.9.

0.0272 $\pm$ 0.0028 $\pm$ 0.0018	<sup>1</sup> ABREU	00R	DLPH	$e^+e^- \rightarrow Z$
0.0192 $\pm$ 0.0025 $\pm$ 0.0005	<sup>2</sup> AKERS	95Q	OPAL	$e^+e^- \rightarrow Z$

- <sup>1</sup> ABREU 00R reports their experiment's uncertainties  $\pm 0.0019 \pm 0.0016 \pm 0.0018$ , where the first error is statistical, the second is systematic, and the third is the uncertainty due to the  $D$  branching fraction. We combine first two in quadrature.  
<sup>2</sup> AKERS 95Q reports  $[\Gamma(\bar{D} \rightarrow D^-\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(D^+ \rightarrow K^-2\pi^+)] = (1.82 \pm 0.20 \pm 0.12) \times 10^{-3}$  which we divide by our best value  $B(D^+ \rightarrow K^-2\pi^+) = (9.46 \pm 0.24) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D^-\pi^-\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0049<math>\pm</math>0.0018<math>\pm</math>0.0007</b>	ABREU	00R	DLPH	$e^+e^- \rightarrow Z$

$\Gamma(D^-\pi^-\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0026<math>\pm</math>0.0015<math>\pm</math>0.0004</b>	ABREU	00R	DLPH	$e^+e^- \rightarrow Z$

$\Gamma(\bar{D}^0\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0681<math>\pm</math>0.0034 OUR AVERAGE</b>				

0.0704 $\pm$ 0.0040 $\pm$ 0.0017	<sup>1</sup> ABREU	00R	DLPH	$e^+e^- \rightarrow Z$
0.064 $\pm$ 0.006 $\pm$ 0.001	<sup>2</sup> AKERS	95Q	OPAL	$e^+e^- \rightarrow Z$

- <sup>1</sup> ABREU 00R reports their experiment's uncertainties  $\pm 0.0034 \pm 0.0036 \pm 0.0017$ , where the first error is statistical, the second is systematic, and the third is the uncertainty due to the  $D$  branching fraction. We combine first two in quadrature.  
<sup>2</sup> AKERS 95Q reports  $[\Gamma(\bar{D} \rightarrow \bar{D}^0\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(D^0 \rightarrow K^-\pi^+)] = (2.52 \pm 0.14 \pm 0.17) \times 10^{-3}$  which we divide by our best value  $B(D^0 \rightarrow K^-\pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{D}^0\pi^-\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0107<math>\pm</math>0.0025<math>\pm</math>0.0011</b>	ABREU	00R	DLPH	$e^+e^- \rightarrow Z$

$\Gamma(\bar{D}^0\pi^+\ell^+\nu_\ell\text{anything})/\Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0023<math>\pm</math>0.0015<math>\pm</math>0.0004</b>	ABREU	00R	DLPH	$e^+e^- \rightarrow Z$











See key on page 601

## Meson Particle Listings

 $B^\pm/B^0/B_s^0/b$ -baryon ADMIXTURE,  $V_{cb}$  and  $V_{ub}$  CKM Matrix Elements

ADAM	96	ZPHY C69 561	W. Adam <i>et al.</i>	(DELPHI Collab.)
ADAM	96D	ZPHY C72 207	W. Adam <i>et al.</i>	(DELPHI Collab.)
ALEXANDER	96	ZPHY C70 357	G. Alexander <i>et al.</i>	(OPAL Collab.)
BUSKULIC	96F	PL B369 151	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	96V	PL B384 471	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	96Y	PL B388 648	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
GROSSMAN	96	NP B465 369	Y. Grossman, Z. Ligeti, E. Nardi	(REHO, CIT)
Also		NP B480 753 (erratum)	Y. Grossman, Z. Ligeti, E. Nardi	
PDG	96	PD 254 1	R. M. Barnett <i>et al.</i>	(PDG Collab.)
UENO	96	PL B381 365	K. Ueno <i>et al.</i>	(AMY Collab.)
ABE.K	95B	PRL 75 3624	K. Abe <i>et al.</i>	(SLD Collab.)
ABREU	95C	PL B347 447	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95D	ZPHY C66 323	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ADAM	95	ZPHY C68 363	W. Adam <i>et al.</i>	(DELPHI Collab.)
AKERS	95Q	ZPHY C67 57	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	95	PL B343 444	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
ABREU	94F	PL B322 459	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	94J	PL B332 488	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	94L	ZPHY C63 3	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	94P	PL B341 109	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	94C	PL B332 201	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	94D	PL B335 542	M. Acciari <i>et al.</i>	(L3 Collab.)
ALBAJAR	94	ZPHY C61 41	C. Albajar <i>et al.</i>	(UA1 Collab.)
BUSKULIC	94G	ZPHY C62 179	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
ABE	93E	PL B313 288	K. Abe <i>et al.</i>	(VENUS Collab.)
ABE	93J	PRL 71 3421	F. Abe <i>et al.</i>	(CDF Collab.)
ABREU	93C	PL B301 145	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	93D	ZPHY C57 181	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	93G	PL B312 253	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	93C	PL B307 247	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	93L	ZPHY C60 217	P.D. Acton <i>et al.</i>	(OPAL Collab.)
ADRIANI	93J	PL B317 467	O. Adriani <i>et al.</i>	(L3 Collab.)
ADRIANI	93K	PL B317 474	O. Adriani <i>et al.</i>	(L3 Collab.)
ADRIANI	93L	PL B317 637	O. Adriani <i>et al.</i>	(L3 Collab.)
AKERS	93B	ZPHY C60 199	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	93B	PL B298 479	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	93O	PL B314 459	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
ABREU	92	ZPHY C53 567	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	92	PL B274 513	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	92C	PL B276 379	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ADEVA	92C	PL B288 395	B. Adeva <i>et al.</i>	(L3 Collab.)
ADRIANI	92	PL B288 412	O. Adriani <i>et al.</i>	(L3 Collab.)
BUSKULIC	92B	PL B284 177	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	92F	PL B295 174	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	92G	PL B295 396	D. Buskalic <i>et al.</i>	(ALEPH Collab.)
ABE	91G	PRL 67 3351	F. Abe <i>et al.</i>	(CDF Collab.)
ADEVA	91C	PL B261 177	B. Adeva <i>et al.</i>	(L3 Collab.)
ADEVA	91H	PL B270 111	B. Adeva <i>et al.</i>	(L3 Collab.)
ALBAJAR	91C	PL B262 163	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALBAJAR	91D	PL B262 171	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALEXANDER	91G	PL B266 485	G. Alexander <i>et al.</i>	(OPAL Collab.)
DECAMP	91C	PL B258 236	D. Decamp <i>et al.</i>	(ALEPH Collab.)
DECAMP	91C	PL B257 492	D. Decamp <i>et al.</i>	(ALEPH Collab.)
ADEVA	90P	PL B252 703	B. Adeva <i>et al.</i>	(L3 Collab.)
BEHREND	90D	ZPHY C47 333	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
HAGEMANN	90	ZPHY C48 401	J. Hagemann <i>et al.</i>	(JADE Collab.)
LYONS	90	PR D41 982	L. Lyons, A.J. Martin, D.H. Saxon	(OXF, BRIS+)
WEIR	90	PL B240 289	A.J. Weir <i>et al.</i>	(Mark II Collab.)
BRAUNSCH...	89B	ZPHY C44 1	R. Braunschweig <i>et al.</i>	(TASSO Collab.)
ONG	89	PRL 62 1236	R.A. Ong <i>et al.</i>	(Mark II Collab.)
BAND	88	PL B200 221	H.R. Band <i>et al.</i>	(MAC Collab.)
KLEM	88	PR D37 41	D.E. Klem <i>et al.</i>	(DELCO Collab.)
ONG	88	PRL 60 2587	R.A. Ong <i>et al.</i>	(Mark II Collab.)
ALBAJAR	87C	PL B186 247	C. Albajar <i>et al.</i>	(UA1 Collab.)
ASH	87	PRL 58 640	W.W. Ash <i>et al.</i>	(MAC Collab.)
BARTEL	87	ZPHY C33 339	W. Bartel <i>et al.</i>	(JADE Collab.)
BROM	87	PL B195 301	J.M. Brom <i>et al.</i>	(HRS Collab.)
PAL	86	PR D33 2708	T. Pal <i>et al.</i>	(DELCO Collab.)
AIHARA	85	ZPHY C27 39	H. Aihara <i>et al.</i>	(TPC Collab.)
BARTEL	85J	PL 163B 277	W. Bartel <i>et al.</i>	(JADE Collab.)
SCHAA	85	PL 160B 188	T. Schaad <i>et al.</i>	(Mark II Collab.)
ALTHOFF	84G	ZPHY C22 219	M. Althoff <i>et al.</i>	(TASSO Collab.)
ALTHOFF	84J	PL 146B 443	M. Althoff <i>et al.</i>	(TASSO Collab.)
KOOP	84	PRL 52 970	D.E. Koop <i>et al.</i>	(DELCO Collab.)
ADEVA	83	PRL 50 799	B. Adeva <i>et al.</i>	(Mark-J Collab.)
ADEVA	83B	PRL 51 443	B. Adeva <i>et al.</i>	(Mark-J Collab.)
BARTEL	83B	PL 132B 241	W. Bartel <i>et al.</i>	(JADE Collab.)
FERNANDEZ	83D	PRL 50 2054	E. Fernandez <i>et al.</i>	(MAC Collab.)
MATTEUZZI	83	PL 129B 141	C. Matteuzzi <i>et al.</i>	(Mark II Collab.)
NELSON	83	PRL 50 1542	M.E. Nelson <i>et al.</i>	(Mark II Collab.)

 $V_{cb}$  and  $V_{ub}$  CKM Matrix Elements

OMITTED FROM SUMMARY TABLE

SEMILEPTONIC BOTTOM HADRON DECAYS AND THE DETERMINATION OF  $V_{cb}$  AND  $V_{ub}$ 

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## INTRODUCTION

Precision determinations of  $|V_{ub}|$  and  $|V_{cb}|$  are central to testing the CKM sector of the Standard Model, and complement the measurements of CP asymmetries in  $B$  decays. The length of the side of the unitarity triangle opposite the well-measured angle  $\beta$  is proportional to the ratio  $|V_{ub}|/|V_{cb}|$ ; its determination is a high priority of the heavy-flavor physics program.

The semileptonic transitions  $b \rightarrow c\ell\bar{\nu}_\ell$  and  $b \rightarrow u\ell\bar{\nu}_\ell$  (where  $\ell$  refers to an electron or muon) each provide two avenues

for determining these CKM matrix elements, namely through inclusive and exclusive final states. Recent measurements and calculations are reflected in the values quoted in this article, which is an update of the previous review [1]. The leptonic decay  $B^- \rightarrow \tau\bar{\nu}$  can also be used to extract  $|V_{ub}|$ ; we do not use this information at present since none of the experimental measurements has reached the significance level of an observation.

The theory underlying the determination of  $|V_{qb}|$  is mature, in particular for  $|V_{cb}|$ . Most of the theoretical approaches use the fact that the mass  $m_b$  of the  $b$  quark is large compared to the scale  $\Lambda_{\text{QCD}}$  that determines low-energy hadronic physics. The basis for precise calculations is a systematic expansion in powers of  $\Lambda/m_b$ , where  $\Lambda \sim 500 - 700$  MeV is a hadronic scale of the order of  $\Lambda_{\text{QCD}}$ , based on effective-field-theory methods described in a separate RPP mini-review [2]. The use of lattice QCD for calculations of non-perturbative quantities plays an essential role in many of the determinations discussed here; lattice methods are discussed in a separate RPP mini-review [3].

The measurements discussed in this review are of branching fractions or ratios of branching fractions. The determination of the  $|V_{cb}|$  and  $|V_{ub}|$  also requires a measurement of the total decay widths of the corresponding  $b$  hadrons, which is the subject of a separate RPP mini-review [4]. The measurements of inclusive semileptonic decays relevant to this review come primarily from  $e^+e^- B$  factories operating at the  $\Upsilon(4S)$  resonance, where  $B\bar{B}$  pairs are produced nearly at rest in the center-of-mass frame. Measurements of exclusive semileptonic decays come from the  $e^+e^- B$  factories and from the LHCb experiment at CERN.

Semileptonic  $B$  meson decay amplitudes to electrons and muons are assumed to be largely free from any impact of non-Standard Model physics, since they are dominated by Standard-Model  $W$  boson exchange. The decays  $\bar{B} \rightarrow D^{(*)}\tau\bar{\nu}_\tau$ , however, provide sensitivity to possible non-universalities in the couplings to the third generation leptons that are present at tree level in models involving new charged mediators. For example, a charged Higgs boson, present in many models of new physics, couples to the mass of the lepton and breaks lepton universality. If the enhanced decay rates seen in recent measurements of these decay modes turn out to be robust, they are an indication of new physics.

Throughout this review the numerical results quoted are based on the methods of the Heavy Flavor Averaging Group [5] using updated values from Ref. 6.

DETERMINATION OF  $|V_{cb}|$ 

*Summary:* The determination of  $|V_{cb}|$  from  $\bar{B} \rightarrow D^*\ell\bar{\nu}_\ell$  decays has a relative precision of about 2%, with comparable contributions from theory and experiment. The value determined from  $\bar{B} \rightarrow D\ell\bar{\nu}_\ell$  decays is consistent and has an uncertainty of 4%. Inclusive decays provide a determination of  $|V_{cb}|$  with a relative uncertainty of about 2%; the limitations

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## $V_{cb}$ and $V_{ub}$ CKM Matrix Elements

arise mainly from our ignorance of higher-order perturbative and non-perturbative corrections.

The values obtained from inclusive and exclusive determinations are only marginally consistent with each other:

$$|V_{cb}| = (42.2 \pm 0.8) \times 10^{-3} \quad (\text{inclusive}) \quad (1)$$

$$|V_{cb}| = (39.2 \pm 0.7) \times 10^{-3} \quad (\text{exclusive}); \quad (2)$$

as a result, their combination should be treated with caution. An average of these determinations has  $p(\chi^2) = 0.33\%$ , so we scale the error by  $\sqrt{\chi^2/1} = 2.9$  to find

$$|V_{cb}| = (40.5 \pm 1.5) \times 10^{-3} \quad (\text{average}). \quad (3)$$

### $|V_{cb}|$ from exclusive decays

Exclusive determinations of  $|V_{cb}|$  make use of semileptonic  $B$  decays into the ground state charmed mesons  $D$  and  $D^*$  and are based on the distribution of the variable  $w \equiv v \cdot v'$ , where  $v$  and  $v'$  are the four velocities of the initial and final-state hadrons. In the rest frame of the decay this variable corresponds to the energy of the final state  $D^{(*)}$  meson. Heavy Quark Symmetry (HQS) [7,8] predicts these decay rates in the infinite mass limit in terms of a single form factor, which is normalized at  $w = 1$ , the point of maximum momentum transfer to the leptons. Measured decay rates and calculations of the form factors are used to determine  $|V_{cb}|$ .

A precise determination requires corrections to the HQS prediction for the normalization as well as some information on the shape of the form factors near the point  $w = 1$ . These calculations utilize Heavy Quark Effective Theory, which is discussed in a separate RPP mini-review [2]. Form factors that are normalized due to HQS are protected against linear corrections [9], and thus the leading corrections are of order  $\Lambda_{\text{QCD}}^2/m_c^2$ . For the form factors that vanish in the infinite mass limit the corrections are in general linear in  $\Lambda_{\text{QCD}}/m_c$ . In addition to these corrections, there are perturbatively calculable radiative corrections from hard gluons and photons, which will be discussed in the relevant sections.

### $\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell$

The decay rate for  $\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell$  is given by

$$\frac{d\Gamma}{dw}(\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell) = \frac{G_F^2 m_B^5}{48\pi^3} |V_{cb}|^2 (w^2 - 1)^{1/2} P(w) (\eta_{\text{ew}} \mathcal{F}(w))^2, \quad (4)$$

where  $P(w)$  is a phase space factor,

$$P(w) = r^3 (1 - r)^2 (w + 1)^2 \left( 1 + \frac{4w}{w + 1} \frac{1 - 2rw + r^2}{(1 - r)^2} \right).$$

with  $r = m_{D^*}/m_B$ . The form factor  $\mathcal{F}(w)$ , which is unity by HQS in the infinite-mass limit, is dominated by the axial vector form factor  $h_{A_1}$  as  $w \rightarrow 1$ . For the definitions of the vector and axial vector form factors as a function of  $w$  see Eq. (2.84) of Ref. 10. The factor  $\eta_{\text{ew}} = 1.015 \pm 0.005$  accounts for the leading electroweak corrections to the four-fermion operator mediating the semileptonic decay [11], and includes an

estimated uncertainty for missing long-distance QED radiative corrections [12].

The determination of  $V_{cb}$  involves an extrapolation to the zero-recoil point. A frequently used one-parameter form for  $\mathcal{F}(w)$  is [13,14]

$$h_{A_1}(w) = \eta_A \left[ 1 + \delta_{1/m^2} + \dots \right] \\ [1 - 8\rho_{A_1}^2 z + (53\rho_{A_1}^2 - 15)z^2 - (231\rho_{A_1}^2 - 91)z^3] \quad (5)$$

with

$$z = (\sqrt{w + 1} - \sqrt{2})/(\sqrt{w + 1} + \sqrt{2}). \quad (6)$$

The use of the variable  $z$  originates from a conformal transformation, which is motivated by analyticity and unitarity. The expansion in this variable converges rapidly in the kinematical region of heavy hadron decays. Expanding in  $w - 1$  one sees that the parameter  $\rho_{A_1}^2$  is the slope of the form factor at  $w = 1$ , and the one parameter form (Eq. (5)) links the curvature (i.e. the coefficient of the second order in the  $w - 1$  expansion) to the slope. All current analyses use the form given in Eq. (5); however, as data become more precise, this simple assumption on the curvature needs to be revised.

The factor  $\eta_A$  is the QCD short-distance radiative correction [15] to the form factor

$$\eta_A = 0.960 \pm 0.007, \quad (7)$$

and  $\delta_{1/m^2}$  comes from non-perturbative  $1/m^2$  corrections, which can be calculated on the lattice (see below).

Precise lattice determinations of the  $B \rightarrow D^{(*)}$  form factors use heavy-quark symmetries, so all uncertainties scale with the deviation of the form factor from unity. The state-of-the-art calculations are “unquenched”, i.e. calculations with realistic sea quarks using  $2 + 1$  flavors. The relevant calculations for the form factor  $\mathcal{F}(w)$  in Ref. 12 quote a total uncertainty at the (1-2)% level. The main contributions to this uncertainty are from the chiral extrapolation from the light quark masses used in the numerical lattice computation to realistic up and down quark masses, and from discretization errors. These sources of uncertainty will be reduced with larger lattice sizes and smaller lattice spacings. Including effects from finite quark masses to calculate the deviation of  $\mathcal{F}(1)$  from unity, the current lattice prediction [12] is

$$\mathcal{F}(1) = 0.906 \pm 0.013, \quad (8)$$

where we take  $\eta_{\text{ew}} = 1.015 \pm 0.005$  [12], appropriate to the mix of  $B^0$  and  $B^+$  decays in the HFAG average, and the errors have been added in quadrature.

Non-lattice estimates based on sum rules for the form factor tend to yield lower values for  $\mathcal{F}(1)$  [16,17,18]. Omitting the contributions from excited states, the sum rules indicate that  $\mathcal{F}(1) < 0.93$ . Including an estimate for the contribution of the excited states yields  $\mathcal{F}(1) = 0.86 \pm 0.01 \pm 0.02$  [18,19] where the second uncertainty accounts for the excited states.

Many experiments [20–28] have measured the differential decay rate as a function of  $w$ , employing a variety of methods:

using either  $B^+$  or  $B^0$  decays, with or without  $B$ -tagging, and with or without explicit reconstruction of the transition pion from  $D^* \rightarrow D$  decays. These measurements are input to a four-dimensional fit [6] for  $\eta_{\text{ew}}\mathcal{F}(1)|V_{cb}|$ ,  $\rho_{A_1}^2$  and the form-factor ratios  $R_1 \propto A_2/A_1$  and  $R_2 \propto V/A_1$ . The fit gives [6]  $\eta_{\text{ew}}\mathcal{F}(1)|V_{cb}| = (35.81 \pm 0.45) \times 10^{-3}$  with a  $p$ -value of 0.15. The leading sources of uncertainty on  $\eta_{\text{ew}}\mathcal{F}(1)|V_{cb}|$  are due to detection efficiencies and  $D^{(*)}$  decay branching fractions.

Along with the lattice value given above for  $\mathcal{F}(1)$  this yields

$$|V_{cb}| = (38.9 \pm 0.5 \pm 0.5 \pm 0.2) \times 10^{-3} \quad (\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell, \text{ LQCD}). \quad (9)$$

where the first error is experimental, the second from lattice QCD and the third from the electroweak and Coulomb correction. The value of  $\mathcal{F}(1)$  obtained from QCD sum rules results in a larger value for  $|V_{cb}|$ :

$$|V_{cb}| = (41.0 \pm 0.5 \pm 1.0) \times 10^{-3} \quad (\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell, \text{ SR}), \quad (10)$$

where the errors are from experiment and theory, respectively.

### $\bar{B} \rightarrow D \ell \bar{\nu}_\ell$

The differential rate for  $\bar{B} \rightarrow D \ell \bar{\nu}_\ell$  is given by

$$\frac{d\Gamma}{dw}(\bar{B} \rightarrow D \ell \bar{\nu}_\ell) = \frac{G_F^2}{48\pi^3} |V_{cb}|^2 (m_B + m_D)^2 m_D^3 (w^2 - 1)^{3/2} (\eta_{\text{ew}} \mathcal{G}(w))^2. \quad (11)$$

The form factor is

$$\mathcal{G}(w) = h_+(w) - \frac{m_B - m_D}{m_B + m_D} h_-(w), \quad (12)$$

where  $h_+$  is normalized to unity due to HQS and  $h_-$  vanishes in the infinite-mass limit. Thus

$$\mathcal{G}(1) = 1 + \mathcal{O}\left(\frac{m_B - m_D}{m_B + m_D} \frac{\Lambda_{\text{QCD}}}{m_c}\right) \quad (13)$$

and the corrections to the HQET predictions are parametrically larger than was the case for  $\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell$ .

Lattice calculations including effects beyond the heavy mass limit have become available, and hence the fact that deviations from the HQET predictions are parametrically larger than in the case  $\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell$  is irrelevant. These unquenched calculations provide information over a range of  $z$  values (see Eq. (6)) and can be used in a simultaneous fit, along with the differential branching fraction, in a form-factor expansion in  $z$  [29,30]. This is important, since the experimental precision near  $w = 1$  is poor given the low decay rate in this region.

From the lattice simulations one obtains the form factor normalization at zero recoil:  $\mathcal{G}(1) = 1.033 \pm 0.095$  [31],  $\mathcal{G}(1) = 1.035 \pm 0.040$  [32] and, from Ref. 33,

$$\mathcal{G}(1) = 1.0528 \pm 0.0082. \quad (14)$$

The most precise measurements of  $\bar{B} \rightarrow D \ell \bar{\nu}_\ell$  [27,34] dominate the average [6] value,  $\eta_{\text{ew}}\mathcal{G}(1)|V_{cb}| = (42.65 \pm 1.53) \times 10^{-3}$ . Using the value from Eq. (14) for  $\mathcal{G}(1)$  and accounting for the

electroweak correction [33],  $\eta_{\text{ew}} = 1.012 \pm 0.005$ , appropriate to the mix of  $B^0$  and  $B^+$  decays in the average, gives

$$|V_{cb}| = (40.0 \pm 1.4 \pm 0.3 \pm 0.2) \times 10^{-3} \quad (\bar{B} \rightarrow D \ell \bar{\nu}_\ell, \text{ LQCD}), \quad (15)$$

where the first uncertainty is from experiment, the second from lattice QCD and the third from the QED and Coulomb corrections.

The first  $|V_{cb}|$  determinations using combined fits to experimental and lattice data over a range of  $q^2$  were reported in Refs. 32 and 33; they find values compatible with the result quoted above. A new preliminary result from Belle [35] provides the most precise single determination,  $\eta_{\text{ew}}|V_{cb}| = (41.10 \pm 1.14) \times 10^{-3}$ . Using the same value for  $\eta_{\text{ew}}$  quoted above results in  $|V_{cb}| = (40.6 \pm 1.1) \times 10^{-3}$ . This new determination has not yet been included in the  $\bar{B} \rightarrow D \ell \bar{\nu}_\ell$  average quoted above.

The  $|V_{cb}|$  averages from  $\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell$  and  $\bar{B} \rightarrow D \ell \bar{\nu}_\ell$  decays are consistent, and their uncertainties are largely uncorrelated. Averaging the results from Eqs. (9) and (15) gives

$$|V_{cb}| = (39.2 \pm 0.7) \times 10^{-3} \quad (\text{exclusive}). \quad (16)$$

### $|V_{cb}|$ from inclusive decays

Measurements of the total semileptonic branching decay rate, along with moments of the lepton energy and hadronic invariant mass spectra in inclusive semileptonic  $b \rightarrow c$  transitions, can be used to determine  $|V_{cb}|$ . The total semileptonic decay rate can be calculated quite reliably in terms of non-perturbative parameters that can be extracted from the information contained in the moments.

### Inclusive semileptonic rate

The theoretical foundation for the calculation of the total semileptonic rate is the Operator Product Expansion (OPE) which yields the Heavy Quark Expansion (HQE) [36,37]. Details can be found in the RPP mini-review on Effective Theories [2].

The OPE result for the total rate can be written schematically (details can be found, *e.g.*, in Ref. 38) as

$$\begin{aligned} \Gamma = & |V_{cb}|^2 \frac{G_F^2 m_b^5(\mu)}{192\pi^3} (1 + A_{\text{ew}}) \times \\ & \left[ z_0^{(0)}(r) + \frac{\alpha_s(\mu)}{\pi} z_0^{(1)}(r) + \left( \frac{\alpha_s(\mu)}{\pi} \right)^2 z_0^{(2)}(r) + \dots \right. \\ & + \frac{\mu_\pi^2}{m_b^2} \left( z_2^{(0)}(r) + \frac{\alpha_s(\mu)}{\pi} z_2^{(1)}(r) + \dots \right) \\ & + \frac{\mu_G^2}{m_b^2} \left( y_2^{(0)}(r) + \frac{\alpha_s(\mu)}{\pi} y_2^{(1)}(r) + \dots \right) \\ & + \frac{\rho_D^3}{m_b^3} \left( z_3^{(0)}(r) + \frac{\alpha_s(\mu)}{\pi} z_3^{(1)}(r) + \dots \right) \\ & \left. + \frac{\rho_{\text{LS}}^3}{m_b^3} \left( y_3^{(0)}(r) + \frac{\alpha_s(\mu)}{\pi} y_3^{(1)}(r) + \dots \right) + \dots \right] \quad (17) \end{aligned}$$

where  $\eta_{\text{ew}} = 1 + A_{\text{ew}}$  denotes the electroweak corrections,  $r$  is the ratio  $m_c/m_b$  and the  $y_i$  and  $z_i$  are functions that appear in the perturbative expansion at different orders of the heavy mass

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expansion. The parameters  $\mu_\pi$ ,  $\mu_G$ ,  $\rho_D$  and  $\rho_{LS}$  constitute the non-perturbative input into the heavy quark expansion; they correspond to certain matrix elements to be discussed below. Similar expansions give the moments of distributions of charged-lepton energy, hadronic invariant mass and hadronic energy, e.g.

$$\langle E_e^n \rangle_{E_e > E_{\text{cut}}} = \int_{E_{\text{cut}}}^{E_{\text{max}}} \frac{d\Gamma}{dE_e} E_e^n dE_e \bigg/ \int_{E_{\text{cut}}}^{E_{\text{max}}} \frac{d\Gamma}{dE_e} dE_e.$$

The OPE result is known up to order  $1/m_b^5$  at tree level [39–42]. The leading term is the parton model, and is known completely to order  $\alpha_s$  and  $\alpha_s^2$  [43–45]; the terms of order  $\alpha_s^{n+1}\beta_0^n$  (where  $\beta_0$  is the first coefficient of the QCD  $\beta$  function,  $\beta_0 = (33 - 2n_f)/3$ ) have been included by the usual BLM procedure [38,46,47]. Corrections of order  $\alpha_s\mu_\pi^2/m_b^2$  have been computed in Refs. 48 and 49, while the  $\alpha_s\mu_G^2/m_b^2$  terms have been calculated in Refs. 50 and 51.

Starting at order  $1/m_b^3$  contributions with an infrared sensitivity to the charm mass,  $m_c$ , appear [41,52,53]. At order  $1/m_b^3$  this “intrinsic charm” contribution manifests as a  $\log(m_c)$  in the coefficient of the Darwin term  $\rho_D^3$ . At higher orders, terms such as  $1/m_b^3 \times 1/m_c^2$  and  $\alpha_s(m_c)1/m_b^3 \times 1/m_c$  appear, which are comparable in size to the contributions of order  $1/m_b^4$ .

The HQE parameters are given in terms of forward matrix elements; the parameters entering the expansion for orders up to  $1/m_b^3$  are ( $D_\perp^\mu = (g_{\mu\nu} - v_\mu v_\nu)D^\nu$ )

$$\begin{aligned} \bar{\Lambda} &= M_B - m_b, \\ \mu_\pi^2 &= -\langle B | \bar{b} (iD_\perp)^2 b | B \rangle, \\ \mu_G^2 &= \langle B | \bar{b} (iD_\perp^\mu) (iD_\perp^\mu) \sigma_{\mu\nu} b | B \rangle, \\ \rho_D^3 &= \langle B | \bar{b} (iD_{\perp\mu}) (ivD) (iD_\perp^\nu) b | B \rangle, \\ \rho_{LS}^3 &= \langle B | \bar{b} (iD_\perp^\mu) (ivD) (iD_\perp^\nu) \sigma_{\mu\nu} b | B \rangle. \end{aligned} \quad (18)$$

These parameters still depend on the heavy quark mass. Sometimes the infinite mass limits of these parameters  $\bar{\Lambda} \rightarrow \bar{\Lambda}_{\text{HQET}}$ ,  $\mu_\pi^2 \rightarrow -\lambda_1$ ,  $\mu_G^2 \rightarrow 3\lambda_2$ ,  $\rho_D^3 \rightarrow \rho_1$  and  $\rho_{LS}^3 \rightarrow 3\rho_2$ , are used instead. The hadronic parameters of the orders  $1/m_b^4$  and  $1/m_b^5$  have been defined and estimated in Ref. 42. The five hadronic parameters  $s_i$  of the order  $1/m_b^4$  can be found in Ref. 40. These terms have not yet been included in the fits.

The rates and the spectra depend strongly on  $m_b$  (or equivalently on  $\bar{\Lambda}$ ). This makes the discussion of renormalization issues mandatory, since the size of QCD corrections is strongly correlated with the definitions used for the quark masses. For example, it is well known that using the pole mass definition for heavy quark masses leads to a perturbative series for the decay rates that does not converge very well, making a precision determination of  $|V_{cb}|$  in such a scheme impossible.

This motivates the use of “short-distance” mass definitions, such as the kinetic scheme [16] or the 1S scheme [54]. Both schemes have been applied to semileptonic  $b \rightarrow c$  transitions and yield comparable results and uncertainties. The 1S scheme eliminates the  $b$  quark pole mass by relating it to the perturbative expression for the mass of the 1S state of the  $\Upsilon$  system.

The physical mass of the  $\Upsilon(1S)$  contains non-perturbative contributions, which have been estimated in Ref. 55. These non-perturbative contributions are small; nevertheless, the best determination of the  $b$  quark mass in the 1S scheme is obtained from sum rules for  $e^+e^- \rightarrow b\bar{b}$  [56]. Alternatively one may use a short-distance mass definition such as the  $\overline{\text{MS}}$  mass,  $m_b^{\overline{\text{MS}}}(m_b)$ . However, it has been argued that the scale  $m_b$  is unnaturally high for  $B$  decays, while for smaller scales  $\mu \sim 1 \text{ GeV}$   $m_b^{\overline{\text{MS}}}(\mu)$  is under poor control. For this reason the so-called “kinetic mass”  $m_b^{\text{kin}}(\mu)$ , has been proposed. It is the mass entering the non-relativistic expression for the kinetic energy of a heavy quark, and is defined using heavy-quark sum rules [16].

### Determination of HQE Parameters and $|V_{cb}|$

Several experiments have measured moments in  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  decays [57–65] as a function of the minimum lepton momentum. The measurements of the moments of the electron energy spectrum (0<sup>th</sup>–3<sup>rd</sup>) and of the squared hadronic mass spectrum (0<sup>th</sup>–2<sup>nd</sup>) have statistical uncertainties that are roughly equal to their systematic uncertainties. The sets of moments measured within each experiment have strong correlations; their use in a global fit requires fully specified statistical and systematic covariance matrices. Measurements of photon energy moments (0<sup>th</sup>–2<sup>nd</sup>) in  $B \rightarrow X_s \gamma$  decays [66–70] as a function of the minimum accepted photon energy are also used in some fits; the dominant uncertainties on these measurements are statistical.

Global fits to the full set of moments [65,67,71–76] have been performed in the 1S and kinetic schemes. The semileptonic moments alone determine a linear combination of  $m_b$  and  $m_c$  very accurately but leave the orthogonal combination poorly determined [77]; additional input is required to allow a precise determination of  $m_b$ . This additional information can come from the radiative  $B \rightarrow X_s \gamma$  moments (with the caveat that the OPE for  $b \rightarrow s \gamma$  breaks down beyond leading order in  $\Lambda_{\text{QCD}}/m_b$ ), which provide complementary information on  $m_b$  and  $\mu_\pi^2$ , or from precise determinations of the charm quark mass [78,79]. The values obtained in the kinetic scheme fits [73,75,76] with these two constraints are consistent. Based on the charm quark mass constraint  $m_c^{\overline{\text{MS}}}(3 \text{ GeV}) = 0.986 \pm 0.013 \text{ GeV}$  [80], a recent analysis [76] obtains

$$|V_{cb}| = (42.21 \pm 0.78) \times 10^{-3} \quad (19)$$

$$m_b^{\text{kin}} = 4.553 \pm 0.020 \text{ GeV} \quad (20)$$

$$\mu_\pi^2(\text{kin}) = 0.465 \pm 0.068 \text{ GeV}^2, \quad (21)$$

where the errors include experimental and theoretical uncertainties.

Theoretical uncertainties are estimated and included in performing the fits. Similar values for the parameters are obtained with a variety of assumptions about the theoretical uncertainties and their correlations. The  $\chi^2/\text{dof}$  is substantially below unity in all fits, suggesting that the theoretical uncertainties may be overestimated. While one could obtain a satisfactory fit with smaller uncertainties, this would result in unrealistically small uncertainties on the extracted HQE parameters, which

are used as input to other calculations (e.g. the determination of  $|V_{ub}|$ ). In any case, the low  $\chi^2$  shows no evidence for duality violations at a significant level. The mass in the  $\overline{\text{MS}}$  scheme corresponding to Eq. (20) is  $m_b^{\overline{\text{MS}}} = 4.18 \pm 0.04 \text{ GeV}$ , which can be compared with a value obtained using relativistic sum rules [80],  $m_b^{\overline{\text{MS}}} = 4.163 \pm 0.016 \text{ GeV}$ , and provides a non-trivial cross-check.

A fit to the measured moments in the 1S scheme [74,5] gives

$$|V_{cb}| = (41.98 \pm 0.45) \times 10^{-3} \quad (22)$$

$$m_b^{1\text{S}} = 4.691 \pm 0.037 \text{ GeV} \quad (23)$$

$$\lambda_1(1\text{S}) = -0.362 \pm 0.067 \text{ GeV}^2, \quad (24)$$

This fit uses semileptonic and radiative moments and constrains the chromomagnetic operator using the  $B^*-B$  and  $D^*-D$  mass differences, but does not include the constraint on  $m_c$  nor the full NNLO corrections.

The fits in the two renormalization schemes give consistent results for  $|V_{cb}|$  and, after translation to a common renormalization scheme, for  $m_b$  and  $\mu_\pi^2$ . We take the fit in the kinetic scheme [76], which includes higher-order corrections and results in a more conservative uncertainty, as the inclusive determination of  $|V_{cb}|$ :

$$|V_{cb}| = (42.2 \pm 0.8) \times 10^{-3} \text{ (inclusive)}. \quad (25)$$

The precision of the global fit results can be further improved by calculating higher-order perturbative corrections to the coefficients of the HQE parameters, in particular the still-missing  $\alpha_s \mu_G^2$  corrections, which are presently only known for  $B \rightarrow X_s \gamma$  [81]. The inclusion of still-higher-order moments, if they can be measured with the required precision, may improve the sensitivity of the fits to higher-order terms in the HQE.

## DETERMINATION OF $|V_{ub}|$

*Summary:* The best determinations of  $|V_{ub}|$  are from  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  decays, where combined fits to theory and experimental data as a function of  $q^2$  provide a precision below 5%; the uncertainties from experiment and theory are comparable in size. Determinations based on inclusive semileptonic decays are done based on different observables and using different calculational ansatzes. All determinations are consistent and provide a precision of about 6%, with comparable contributions to the uncertainty from experiment and theory.

The values obtained from inclusive and exclusive determinations are

$$|V_{ub}| = (4.49 \pm 0.16 \pm_{0.18}^{0.16}) \times 10^{-3} \text{ (inclusive)}, \quad (26)$$

$$|V_{ub}| = (3.72 \pm 0.19) \times 10^{-3} \text{ (exclusive)}. \quad (27)$$

The two determinations are independent, and the dominant uncertainties are on multiplicative factors. To combine these values, the inclusive and exclusive values are weighted by their relative errors and the uncertainties are treated as normally

distributed. The resulting average has  $p(\chi^2) = 1.0\%$ , so we scale the error by  $\sqrt{\chi^2/1} = 2.6$  to find

$$|V_{ub}| = (4.09 \pm 0.39) \times 10^{-3} \text{ (average)}. \quad (28)$$

Given the poor consistency between the two determinations, this average should be treated with caution.

### $|V_{ub}|$ from inclusive decays

The theoretical description of inclusive  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays is based on the Heavy Quark Expansion, as for  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  decays, and leads to a predicted total decay rate with uncertainties below 5% [82,83]. Unfortunately, the total decay rate is hard to measure due to the large background from CKM-favored  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  transitions. Technically, the calculation of the partial decay rate in regions of phase space where  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  decays are suppressed requires the introduction of a non-perturbative distribution function, the “shape function” (SF) [84,85], whose form is unknown. The shape function becomes important when the light-cone momentum component  $P_+ \equiv E_X - |P_X|$  is not large compared to  $\Lambda_{QCD}$ , as is the case near the endpoint of the  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  lepton spectrum. Partial rates for  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  are predicted and measured in a variety of kinematic regions that differ in their sensitivity to shape-function effects.

At leading order a single shape function appears, which is universal for all heavy-to-light transitions [84,85] and can be measured in  $\bar{B} \rightarrow X_s \gamma$  decays. At subleading order in  $1/m_b$ , several shape functions appear [86]. Thus, prescriptions that relate directly the partial rates for  $\bar{B} \rightarrow X_s \gamma$  and  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays [87–90] are limited to leading order in  $1/m_b$ .

Existing approaches have tended to use parameterizations of the leading SF that respect constraints on the normalization and on the first and second moments, which are given in terms of the HQE parameters  $\bar{\Lambda} = M_B - m_b$  and  $\mu_\pi^2$ , respectively. The relations between SF moments and HQE parameters are known to second order in  $\alpha_s$  [91]. As a result, measurements of HQE parameters from global fits to  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  and  $\bar{B} \rightarrow X_s \gamma$  moments can be used to constrain the SF moments, as well as to provide accurate values of  $m_b$  and other parameters for use in determining  $|V_{ub}|$ . The authors of Ref. 92 propose the use of a set of orthogonal basis functions to approximate the SF and thereby include the known short-distance contributions and renormalization properties of the SF; this would allow a global fit of all inclusive  $B$  meson decay data.

The calculations used for the fits performed by HFAG are documented in Ref. 93 (BLNP), Ref. 94 (GGOU), Ref. 95 (DGE) and Ref. 96 (BLL).

The triple differential rate in the variables

$$P_l = M_B - 2E_l, \quad P_- = E_X + |\vec{P}_X|, \quad P_+ = E_X - |\vec{P}_X| \quad (29)$$

is

$$\frac{d^3\Gamma}{dP_+ dP_- dP_l} = \frac{G_F^2 |V_{ub}|^2}{16\pi^2} (M_B - P_+) \left\{ (P_- - P_l)(M_B - P_- + P_l - P_+) \mathcal{F}_1 \right. \\ \left. + (M_B - P_-)(P_- - P_+) \mathcal{F}_2 + (P_- - P_l)(P_l - P_+) \mathcal{F}_3 \right\}. \quad (30)$$

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The “structure functions”  $\mathcal{F}_i$  can be calculated using factorization theorems that have been proven to subleading order in the  $1/m_b$  expansion.

The BLNP [93] calculation uses these factorization theorems to write the  $\mathcal{F}_i$  in terms of perturbatively calculable hard coefficients  $H$  and jet functions  $J$ , which are convolved with the (soft) light-cone distribution functions  $S$ , the shape functions of the  $B$  meson. The calculation of  $\mathcal{O}(\alpha_s^2)$  contributions [97,98] is not yet complete and is not included in the  $|V_{ub}|$  determination given below.

The leading order term in the  $1/m_b$  expansion of the  $\mathcal{F}_i$  contains a single non-perturbative function and is calculated to subleading order in  $\alpha_s$ , while at subleading order in the  $1/m_b$  expansion there are several independent non-perturbative functions that have been calculated only at tree level in the  $\alpha_s$  expansion.

A distinct approach (GGOU) [94] uses a hard, Wilsonian cut-off that matches the definition of the kinetic mass. The non-perturbative input is similar to what is used in BLNP, but the shape functions are defined differently. In particular, they are defined at finite  $m_b$  and depend on the light-cone component  $k_+$  of the  $b$  quark momentum and on the momentum transfer  $q^2$  to the leptons. These functions include subleading effects to all orders; as a result they are non-universal, with one shape function corresponding to each structure function in Eq. (30). Their  $k_+$  moments can be computed in the OPE and related to observables and to the shape functions defined in Ref. 93.

Going to subleading order in  $\alpha_s$  requires the definition of a renormalization scheme for the HQE parameters and for the SF. The relation between the moments of the SF and the forward matrix elements of local operators is plagued by ultraviolet problems and requires additional renormalization. A scheme for improving this behavior was suggested in Refs. 93 and 99, which introduce a definition of the quark mass (the so-called shape-function scheme) based on the first moment of the measured  $\bar{B} \rightarrow X_s \gamma$  photon energy spectrum. Likewise, the HQE parameters can be defined from measured moments of spectra, corresponding to moments of the SF.

One can attempt to calculate the SF by using additional assumptions. One approach (DGE) is the so-called “dressed gluon exponentiation” [95], where the perturbative result is continued into the infrared regime using the renormalon structure obtained in the large  $\beta_0$  limit, where  $\beta_0$  has been defined following Eq. (17).

In order to reduce sensitivity to SF uncertainties, measurements that use a combination of cuts on the leptonic momentum transfer  $q^2$  and the hadronic invariant mass  $m_X$ , as suggested in Ref. 96, have been made. In general, efforts to extend the experimental measurements of  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  into charm-dominated regions (in order to reduce SF uncertainties) lead to an increased experimental sensitivity to the modeling of  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays, resulting in measured partial rates with an undesirable level of model dependence. The measurements quoted below have used

a variety of functional forms to parameterize the leading SF; in no case does this lead to more than a 2% uncertainty on  $|V_{ub}|$ .

Weak Annihilation [100,101,94] (WA) can in principle contribute significantly in the high- $q^2$  region of  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays. Estimates based on semileptonic  $D_s$  decays [101,53,96] lead to a  $\sim 2\%$  uncertainty on the total  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  rate from the  $\Upsilon(4S)$ . The  $q^2$  spectrum of the WA contribution is not well known, but from the OPE it is expected to contribute predominantly at high  $q^2$ . More recent investigations [53,102,103] indicate that WA is a small effect, but may become a significant source of uncertainty for  $|V_{ub}|$  measurements that accept only a small fraction of the full  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  phase space.

### Measurements

We summarize the measurements used in the determination of  $|V_{ub}|$  below. Given the improved precision and more rigorous theoretical interpretation of the recent measurements, earlier determinations [104–107] will not be considered in this review.

Inclusive electron momentum measurements [108–110] reconstruct a single charged electron to determine a partial decay rate for  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  near the kinematic endpoint. This results in a selection efficiency of order 50% and only modest sensitivity to the modeling of detector response. The inclusive electron momentum spectrum from  $B\bar{B}$  events, after subtraction of the  $e^+e^- \rightarrow q\bar{q}$  continuum background, is fitted to a model  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  spectrum and several components ( $D\ell\bar{\nu}_\ell$ ,  $D^*\ell\bar{\nu}_\ell$ , ...) of the  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  background; the dominant uncertainties are related to this subtraction and modelling. The decay rate can be cleanly extracted for  $E_e > 2.3$  GeV, but this is deep in the SF region, where theoretical uncertainties are large. Measurements with  $E_e > 2.0$  GeV have a low ( $< 1/10$ ) but usable signal-to-background (S/B) ratio. The resulting  $|V_{ub}|$  values for various  $E_e$  cuts are given in Table 1.

An untagged “neutrino reconstruction” measurement [111] from BABAR uses a combination [112] of a high-energy electron with a measurement of the missing momentum vector. This allows S/B  $\sim 0.7$  for  $E_e > 2.0$  GeV and a  $\approx 5\%$  selection efficiency, but at the cost of a smaller accepted phase space for  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays and uncertainties associated with the determination of the missing momentum. The corresponding values for  $|V_{ub}|$  are given in Table 1.

The large samples accumulated at the  $B$  factories allow studies in which one  $B$  meson is fully reconstructed and the recoiling  $B$  decays semileptonically [113–117]. The experiments can fully reconstruct a “tag”  $B$  candidate in about 0.5% (0.3%) of  $B^+B^-$  ( $B^0\bar{B}^0$ ) events. An electron or muon with center-of-mass momentum above 1.0 GeV is required amongst the charged tracks not assigned to the tag  $B$  and the remaining particles are assigned to the  $X_u$  system. The full set of kinematic properties ( $E_\ell$ ,  $m_X$ ,  $q^2$ , etc.) are available for studying the semileptonically decaying  $B$ , making possible selections that accept up to 90% of the full  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  rate. Despite requirements (e.g. on the square of the missing mass) aimed at rejecting events with additional missing particles, undetected

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or mis-measured particles from  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  decay (e.g.,  $K_L^0$  and additional neutrinos) remain an important source of uncertainty. Measurements with the largest kinematic acceptance (i.e.  $E_\ell > 1 \text{ GeV}$ ) lead to the smallest overall uncertainties on  $|V_{ub}|$ .

BABAR [113] and Belle [114,115] have measured partial rates with cuts on  $m_X$ ,  $m_X$  and  $q^2$ ,  $P_+$  and  $E_\ell$  using the recoil method. In each case the experimental systematics have significant contributions from the modeling of  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  and  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  decays and from the detector response to charged particles, photons and neutral hadrons. The corresponding  $|V_{ub}|$  values are given in Table 1.

#### $|V_{ub}|$ from inclusive partial rates

The measured partial rates and theoretical calculations from BLNP, GGOU and DGE described previously are used to determine  $|V_{ub}|$  from all measured partial  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  rates [6]; selected values are given in Table 1. The correlations amongst the multiple BABAR recoil-based measurements [113] are fully accounted for in the average. The statistical correlations amongst the other measurements used in the average are tiny (due to small overlaps among signal events and large differences in S/B ratios) and have been ignored. Correlated systematic and theoretical errors are taken into account, both within an experiment and between experiments. As an illustration of the relative sizes of the uncertainties entering  $|V_{ub}|$  we give the error breakdown for the GGOU average: statistical—1.9%; experimental—1.7%;  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  modeling—1.3%;  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  modeling—1.9%; HQE parameters ( $m_b$ ) —1.6%; higher-order corrections—1.5%;  $q^2$  modeling—1.4%; Weak Annihilation— $^{+0.0\%}_{-2.0\%}$ ; SF parameterization—0.2%.

The averages quoted here are based on the following  $m_b$  values:  $m_b^{SF} = 4.569 \pm 0.029 \text{ GeV}$  for BLNP,  $m_b^{\text{kin}} = 4.541 \pm 0.023 \text{ GeV}$  for GGOU, and  $m_b^{MS} = 4.177 \pm 0.043 \text{ GeV}$  for DGE. The  $m_b^{\text{kin}}$  value is determined in a global fit to moments in the kinetic scheme; this value is translated into  $m_b^{SF}$  and  $m_b^{MS}$  at fixed order in  $\alpha_s$ . These input values are based on an earlier determination of  $m_b^{\text{kin}}$  than is quoted in equation Eq. (20); using the latest value would decrease the  $|V_{ub}|$  averages by 1-2%.

The theoretical calculations produce very similar results for  $|V_{ub}|$ ; the standard deviation of the  $|V_{ub}|$  values for the  $E_\ell > 2.1 \text{ GeV}$  rate is 4.9%, for the  $m_X - q^2$  rate is 0.6%, and for the  $E_\ell > 1 \text{ GeV}$  rate is 1.6%. The  $|V_{ub}|$  values do not show a marked trend versus the kinematic acceptance,  $f_u$ , for  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays. The  $p$ -values of the averages are in the range 55-62%, indicating that the ratios of calculated partial widths in the different phase space regions are in good agreement with ratios of measured partial branching fractions.

Hadronization uncertainties also impact the  $|V_{ub}|$  determination. The theoretical expressions are valid at the parton level and do not incorporate any resonant structure (e.g.  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$ ); this must be added to the simulated  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  event samples, since the detailed final state multiplicity and structure impacts the estimates of experimental acceptance and efficiency. The

Ref.	cut (GeV)	BLNP	GGOU	DGE
[108]	$E_e > 2.1$	$428 \pm 50 \begin{smallmatrix} +31 \\ -36 \end{smallmatrix}$	$421 \pm 49 \begin{smallmatrix} +23 \\ -33 \end{smallmatrix}$	$390 \pm 45 \begin{smallmatrix} +26 \\ -28 \end{smallmatrix}$
[111]	$E_e - q^2$	$453 \pm 22 \begin{smallmatrix} +33 \\ -38 \end{smallmatrix}$	not available	$417 \pm 20 \begin{smallmatrix} +28 \\ -29 \end{smallmatrix}$
[110]	$E_e > 2.0$	$454 \pm 26 \begin{smallmatrix} +27 \\ -33 \end{smallmatrix}$	$450 \pm 26 \begin{smallmatrix} +18 \\ -25 \end{smallmatrix}$	$434 \pm 25 \begin{smallmatrix} +23 \\ -25 \end{smallmatrix}$
[109]	$E_e > 1.9$	$493 \pm 46 \begin{smallmatrix} +27 \\ -29 \end{smallmatrix}$	$493 \pm 46 \begin{smallmatrix} +17 \\ -22 \end{smallmatrix}$	$485 \pm 45 \begin{smallmatrix} +21 \\ -25 \end{smallmatrix}$
[113]	$q^2 > 8$	$430 \pm 23 \begin{smallmatrix} +26 \\ -28 \end{smallmatrix}$	$432 \pm 23 \begin{smallmatrix} +27 \\ -30 \end{smallmatrix}$	$427 \pm 22 \begin{smallmatrix} +20 \\ -20 \end{smallmatrix}$
[113]	$m_X < 1.7$	$415 \pm 25 \begin{smallmatrix} +28 \\ -27 \end{smallmatrix}$	$424 \pm 26 \begin{smallmatrix} +32 \\ -32 \end{smallmatrix}$	$424 \pm 26 \begin{smallmatrix} +37 \\ -32 \end{smallmatrix}$
[113]	$P_+ < 0.66$	$430 \pm 20 \begin{smallmatrix} +28 \\ -27 \end{smallmatrix}$	$429 \pm 20 \begin{smallmatrix} +31 \\ -22 \end{smallmatrix}$	$453 \pm 21 \begin{smallmatrix} +24 \\ -22 \end{smallmatrix}$
[113]	$m_X < 1.55$	$430 \pm 20 \begin{smallmatrix} +28 \\ -27 \end{smallmatrix}$	$429 \pm 20 \begin{smallmatrix} +31 \\ -22 \end{smallmatrix}$	$453 \pm 21 \begin{smallmatrix} +24 \\ -22 \end{smallmatrix}$
[113]	$E_\ell > 1$	$432 \pm 24 \begin{smallmatrix} +19 \\ -21 \end{smallmatrix}$	$442 \pm 24 \begin{smallmatrix} +9 \\ -11 \end{smallmatrix}$	$446 \pm 24 \begin{smallmatrix} +13 \\ -13 \end{smallmatrix}$
[115]	$E_\ell > 1$	$449 \pm 27 \begin{smallmatrix} +20 \\ -22 \end{smallmatrix}$	$460 \pm 27 \begin{smallmatrix} +10 \\ -11 \end{smallmatrix}$	$463 \pm 28 \begin{smallmatrix} +13 \\ -13 \end{smallmatrix}$
HFAG average		$445 \pm 16 \begin{smallmatrix} +21 \\ -22 \end{smallmatrix}$	$451 \pm 16 \begin{smallmatrix} +12 \\ -15 \end{smallmatrix}$	$452 \pm 16 \begin{smallmatrix} +15 \\ -16 \end{smallmatrix}$

**Table 1:**  $|V_{ub}|$  (in units of  $10^{-5}$ ) from inclusive  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  measurements. The first uncertainty on  $|V_{ub}|$  is experimental, while the second includes both theoretical and HQE parameter uncertainties. The values are listed in order of increasing kinematic acceptance  $f_u$  (0.19 to 0.90); those below the horizontal bar are based on recoil methods.

experiments have adopted procedures to input resonant structure while preserving the appropriate behavior in the kinematic variables ( $q^2, E_\ell, m_X$ ) averaged over the sample, but these prescriptions are *ad hoc*. The resulting uncertainties have been estimated to be  $\sim 1\text{-}2\%$  on  $|V_{ub}|$ .

All calculations yield compatible  $|V_{ub}|$  values and similar error estimates. We take the arithmetic mean of the values and errors to find

$$|V_{ub}| = (4.49 \pm 0.16_{\text{exp}} \begin{smallmatrix} +0.16 \\ -0.18 \end{smallmatrix}_{\text{theo}}) \times 10^{-3} \quad (\text{inclusive}). \quad (31)$$

The BLL [96] calculation has been used along with measurements [113,114,118] with cuts on  $m_X$  and  $q^2$  to determine  $|V_{ub}|$ . Using  $m_b^{\text{IS}} = 4.704 \pm 0.029 \text{ GeV}$  yields a  $|V_{ub}|$  value of  $(4.62 \pm 0.20 \pm 0.29) \times 10^{-3}$ , which is somewhat larger than the corresponding values listed in Table 1.

#### $|V_{ub}|$ from exclusive decays

Exclusive charmless semileptonic decays offer a complementary means of determining  $|V_{ub}|$ . For the experiments, the specification of the final state provides better background rejection, but the branching fraction to a specific final state is typically only a few percent of that for inclusive decays. For theory, the calculation of the form factors for  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays is challenging, but brings in a different set of uncertainties from those encountered in inclusive decays. In this review we focus on  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$ , as it is the most promising mode for both experiment and theory. Measurements of other exclusive  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  states can be found in Refs. [119–126].

#### $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$ form factor calculations

The relevant form factors for the decay  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  are usually defined as

$$\langle \pi(p_\pi) | V^\mu | B(p_B) \rangle = \quad (32)$$

$$f_+(q^2) \left[ p_B^\mu + p_\pi^\mu - \frac{m_B^2 - m_\pi^2}{q^2} q^\mu \right] + f_0(q^2) \frac{m_B^2 - m_\pi^2}{q^2} q^\mu$$

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in terms of which the rate becomes (in the limit  $m_\ell \rightarrow 0$ )

$$\frac{d\Gamma}{dq^2} = \frac{G_F^2 |V_{ub}|^2}{24\pi^3} |p_\pi|^3 |f_+(q^2)|^2, \quad (33)$$

where  $p_\pi$  is the pion momentum in the  $B$  meson rest frame.

Currently available non-perturbative methods for the calculation of the form factors include lattice QCD (LQCD) and light-cone sum rules (LCSR). The two methods are complementary in phase space, since the lattice calculation is restricted to the kinematical range of high momentum transfer  $q^2$  to the leptons, while light-cone sum rules provide information near  $q^2 = 0$ . Interpolations between these two regions can be constrained by unitarity and analyticity.

Unquenched simulations, i.e. where quark loop effects are fully incorporated, are now standard, and have been performed by the Fermilab/MILC [127], the HPQCD [128] and the RBC/UKQCD [129] collaborations. The calculations differ in the way the  $b$  quark is simulated, with HPQCD using nonrelativistic QCD, and Fermilab/MILC and RBC/UKQCD using relativistic  $b$  quarks with the Fermilab and Columbia heavy-quark formulations; they agree within the quoted errors. The result from Ref. 127 represents a significant improvement in precision. The form factor  $f_+$  evaluated at  $q^2 = 20 \text{ GeV}^2$  has an estimated uncertainty of 3.4%, where the leading contribution is due to the chiral-continuum extrapolation fit, which includes statistical and heavy-quark discretization errors. However, the lattice simulations are restricted to the region of large  $q^2$ , i.e. the region  $q_{\text{max}}^2 > q^2 \gtrsim 15 \text{ GeV}^2$ .

The extrapolation to small values of  $q^2$  is performed using guidance from analyticity and unitarity. Making use of the heavy-quark limit, stringent constraints on the shape of the form factor can be derived [130], and the conformal mapping of the kinematical variables onto the complex unit disc yields a rapidly converging series in the variable

$$z = \frac{\sqrt{t_+ - t_-} - \sqrt{t_+ - q^2}}{\sqrt{t_+ - t_-} + \sqrt{t_+ - q^2}}$$

where  $t_\pm = (M_B \pm m_\pi)^2$ . The use of lattice data in combination with experimental measurements of the differential decay rate provides a stringent constraint on the shape of the form factor in addition to precise determination of  $|V_{ub}|$  [131].

Another established non-perturbative approach to obtain the form factors is through Light-Cone QCD Sum Rules (LCSR). The sum-rule approach provides an estimate for the product  $f_B f_+(q^2)$ , valid in the region  $0 < q^2 \lesssim 12 \text{ GeV}^2$ . The determination of  $f_+(q^2)$  itself requires knowledge of the decay constant  $f_B$ , which is usually obtained by replacing  $f_B$  by its two-point QCD (SVZ) sum rule [132] in terms of perturbative and condensate contributions. The advantage of this procedure is the approximate cancellation of various theoretical uncertainties in the ratio  $(f_B f_+)/f_B$ .

The LCSR for  $f_B f_+$  is based on the light-cone OPE of the relevant vacuum-to-pion correlation function, calculated in full QCD at finite  $b$ -quark mass. The resulting expressions actually

**Table 2:** Total and partial branching fractions for  $\bar{B}^0 \rightarrow \pi^+ \ell^- \bar{\nu}_\ell$ .  $B$ -tagged analyses are indicated (SL for *semileptonic*, Had for *hadronic*). The first uncertainty listed is from statistics, the second from systematics. Measurements of  $\mathcal{B}(B^- \rightarrow \pi^0 \ell^- \bar{\nu}_\ell)$  have been multiplied by a factor  $2\tau_{B^0}/\tau_{B^+}$  to obtain the values below.

	$\mathcal{B} \times 10^4$	$\mathcal{B}(q^2 > 16) \times 10^4$
CLEO $\pi^+, \pi^0$ [124]	$1.38 \pm 0.15 \pm 0.11$	$0.41 \pm 0.08 \pm 0.04$
BABAR $\pi^+, \pi^0$ [125]	$1.41 \pm 0.05 \pm 0.08$	$0.32 \pm 0.02 \pm 0.03$
BABAR $\pi^+$ [126]	$1.44 \pm 0.04 \pm 0.06$	$0.37 \pm 0.02 \pm 0.02$
Belle $\pi^+, \pi^0$ [139]	$1.48 \pm 0.04 \pm 0.07$	$0.40 \pm 0.02 \pm 0.02$
Belle SL $\pi^+$ [140]	$1.41 \pm 0.19 \pm 0.15$	$0.37 \pm 0.10 \pm 0.04$
Belle SL $\pi^0$ [140]	$1.41 \pm 0.26 \pm 0.15$	$0.37 \pm 0.15 \pm 0.04$
Belle Had $\pi^+$ [119]	$1.49 \pm 0.09 \pm 0.07$	$0.45 \pm 0.05 \pm 0.02$
Belle Had $\pi^0$ [119]	$1.48 \pm 0.15 \pm 0.08$	$0.36 \pm 0.07 \pm 0.02$
BABAR SL $\pi^+$ [141]	$1.38 \pm 0.21 \pm 0.08$	$0.46 \pm 0.13 \pm 0.03$
BABAR SL $\pi^0$ [141]	$1.78 \pm 0.28 \pm 0.15$	$0.44 \pm 0.17 \pm 0.06$
BABAR Had $\pi^+$ [142]	$1.07 \pm 0.27 \pm 0.19$	$0.65 \pm 0.20 \pm 0.13$
BABAR Had $\pi^0$ [142]	$1.52 \pm 0.41 \pm 0.30$	$0.48 \pm 0.22 \pm 0.12$
Average	$1.45 \pm 0.02 \pm 0.04$	$0.38 \pm 0.01 \pm 0.01$

comprise a triple expansion: in the twist  $t$  of the operators near the light-cone, in  $\alpha_s$ , and in the deviation of the pion distribution amplitudes from their asymptotic form, which is fixed from conformal symmetry. The sources of uncertainties in the LCSR calculation are discussed in Refs. 133 and 134; currently the total uncertainty slightly larger than 10% on  $|V_{ub}|$  is obtained from a LCSR calculation of

$$\begin{aligned} \Delta\zeta(0, q_{\text{max}}^2) &= \frac{G_F^2}{24\pi^3} \int_0^{q_{\text{max}}^2} dq^2 p_\pi^3 |f_+(q^2)|^2 \\ &= \frac{1}{|V_{ub}|^2 \tau_{B^0}} \int_0^{q_{\text{max}}^2} dq^2 \frac{d\mathcal{B}(B \rightarrow \pi \ell \nu)}{dq^2} \end{aligned} \quad (34)$$

which gives [135]

$$\Delta\zeta(0, 12 \text{ GeV}^2) = 4.59_{-0.85}^{+1.00} \text{ ps}^{-1}. \quad (35)$$

The recent calculation of two loop contributions to the LCQCD sum rules [136] and the estimation of statistical correlations [137] results in only small changes to the central value and uncertainty.

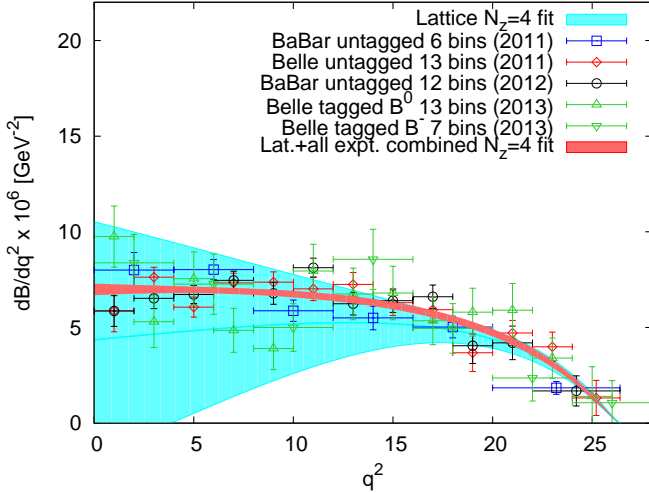
### $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$ measurements

The  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  measurements fall into two broad classes: untagged, in which case the reconstruction of the missing momentum of the event serves as an estimator for the unseen neutrino, and tagged, in which the second  $B$  meson in the event is fully reconstructed in either a hadronic or semileptonic decay mode. The tagged measurements have high and uniform acceptance and S/B as high as 10, but low statistics. The untagged measurements have somewhat higher background levels ( $S/B < 1$ ) and make slightly more restrictive kinematic cuts, but provide better precision on the  $q^2$  dependence of the form factor.



CLEO has analyzed  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  and  $\bar{B} \rightarrow \rho \ell \bar{\nu}_\ell$  using an untagged analysis [124]. Similar analyses have been done at BABAR [125,126] and Belle [139]. The leading systematic uncertainties in the untagged  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  analyses are associated with modeling the missing momentum reconstruction, with backgrounds from  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays and  $e^+e^- \rightarrow q\bar{q}$  continuum events, and with varying the form factor for the  $\bar{B} \rightarrow \rho \ell \bar{\nu}_\ell$  decay. The values obtained for the full and partial branching fractions [6] are listed in Table 2 above the horizontal line. These BABAR and Belle measurements provide the differential  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  rate versus  $q^2$ , shown in Fig. 1, which is used in the determination of  $|V_{ub}|$  discussed below.

Analyses [140,141] based on reconstructing a  $B$  in the  $\bar{D}^{(*)} \ell^+ \nu_\ell$  decay mode and looking for a  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  or  $\bar{B} \rightarrow \rho \ell \bar{\nu}_\ell$  decay amongst the remaining particles in the event make use of the fact that the  $B$  and  $\bar{B}$  are back-to-back in the  $\Upsilon(4S)$  frame to construct a discriminant variable that provides a signal-to-noise ratio above unity for all  $q^2$  bins. A related technique was discussed in Ref. 143. BABAR [141] and Belle [119] have also used their samples of  $B$  mesons reconstructed in hadronic decay modes to measure exclusive charmless semileptonic decays, resulting in very clean but small samples. The corresponding full and partial branching fractions are given in Table 2. The averages take account of correlations and common systematic uncertainties, and have  $p(\chi^2) > 0.5$  in each case.



**Figure 1:** The differential  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  branching fraction versus  $q^2$  for untagged measurements along with a combined fit to lattice calculations from Ref. 127.

$|V_{ub}|$  can be obtained from the average  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  branching fraction and the measured  $q^2$  spectrum. Fits to the  $q^2$  spectrum using a theoretically motivated parameterization (e.g. "BCL" from Ref. 138) remove most of the model dependence from theoretical uncertainties in the shape of the spectrum. The most sensitive method for determining  $|V_{ub}|$  from  $\bar{B} \rightarrow \pi \ell \bar{\nu}_\ell$  decays employs a simultaneous fit [6,127,144] to measured

experimental partial rates and lattice points versus  $q^2$  (or  $z$ ) to determine  $|V_{ub}|$  and the first few coefficients of the expansion of the form factor in  $z$ . The most precise determination at present is from Ref. 127, which uses as experimental input the measurements in Refs. [119,125,126,139], and finds  $|V_{ub}| = (3.72 \pm 0.16) \times 10^{-3}$ . This fit, shown in Fig. 1, has  $p(\chi^2) = 2\%$ . We scale the quoted uncertainty by  $\sqrt{\chi^2/\text{d.o.f.}} = 1.2$  to find

$$|V_{ub}| = (3.72 \pm 0.19) \times 10^{-3} \quad (\text{exclusive}). \quad (36)$$

We note that fits that exclude the data of Ref. 125 have larger  $p(\chi^2)$  and result in higher  $|V_{ub}|$  values between  $3.8 \times 10^{-3}$  and  $4.0 \times 10^{-3}$ . The contributions to the uncertainty from the lattice calculation and from the experimental measurements are comparable.

An alternative approach using the average [6] of partial branching fractions in the  $q^2 < 12 \text{ GeV}^2$  region,  $(0.81 \pm 0.02 \pm 0.02) \times 10^{-4}$ , along with an LCSR calculation of the theoretical rate [135], gives

$$|V_{ub}| = (3.41 \pm 0.06_{\text{exp}}^{+0.37}_{-0.32} \text{theo}) \times 10^{-3} \quad (\text{LCSR}, q^2 < 12 \text{ GeV}^2). \quad (37)$$

#### SEMILEPTONIC $B$ -BARYON DECAYS AND DETERMINATION OF $|V_{ub}|/|V_{cb}|$

*Summary:* A significant sample of  $\Lambda_b^0$  baryons is available at the LHCb experiment, and methods have been developed to study their semileptonic decays. Both  $\Lambda_b^0 \rightarrow p \mu \bar{\nu}$  and  $\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \bar{\nu}$  decays have been measured at LHCb, and the ratio of branching fractions to these two decay modes is used to determine the ratio  $|V_{ub}|/|V_{cb}|$ . Averaging the LHCb determination with those obtained from inclusive and exclusive  $B$  meson decays, we find

$$|V_{ub}|/|V_{cb}| = 0.096 \pm 0.007 \quad (\text{average})$$

where the uncertainty has been scaled by a factor  $\sqrt{\chi^2/\text{ndf}} = 2.0$ . In light of the poor consistency of the three determinations considered, the average should be treated with caution.

#### $\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \bar{\nu}$ and $\Lambda_b^0 \rightarrow p \mu \bar{\nu}$

The  $\Lambda_b^0 \rightarrow \Lambda_c^+$  and  $\Lambda_b^0 \rightarrow p$  semileptonic transitions are described in terms of six form factors each. The three form factors corresponding to the vector current can be defined as [145]

$$\begin{aligned} \langle F(p', s') | \bar{q} \gamma_\mu b | \Lambda_b^0(p, s) \rangle = & \bar{u}_F(p', s') \left\{ f_0(q^2) (M_{\Lambda_b^0} - m_F) \frac{q_\mu}{q^2} \right. \\ & + f_+(q^2) \frac{M_{\Lambda_b^0} + m_F}{s_+} \left( p_\mu + p'_\mu - \frac{q_\mu}{q^2} (M_{\Lambda_b^0}^2 - m_F^2) \right) \\ & \left. + f_\perp(q^2) \left( \gamma_\mu - \frac{2m_F}{s_+} p_\mu - \frac{2M_{\Lambda_b^0}}{s_+} p'_\mu \right) \right\} u_{\Lambda_b^0}(p, s), \quad (38) \end{aligned}$$

where  $F = p$  or  $\Lambda_c^+$  and where we define  $s_\pm = (M_{\Lambda_b^0} \pm m_F)^2 - q^2$ . At vanishing momentum transfer,  $q^2 \rightarrow 0$ , the kinematic constraint  $f_0(0) = f_+(0)$  holds. The form factors are defined in such a way that they correspond to time-like

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(scalar), longitudinal and transverse polarization with respect to the momentum-transfer  $q^\mu$  for  $f_0$ ,  $f_+$  and  $f_\perp$ , respectively. Furthermore we have chosen the normalization in such a way that for  $f_0, f_+, f_\perp \rightarrow 1$  one recovers the expression for point-like baryons.

Likewise, the expression for the axial-vector current is

$$\begin{aligned} \langle F(p', s') | \bar{q} \gamma_\mu \gamma_5 b | \Lambda_b^0(p, s) \rangle = & -\bar{u}_F(p', s') \gamma_5 \\ & \left\{ g_0(q^2) (M_{\Lambda_b^0} + m_F) \frac{q_\mu}{q^2} \right. \\ & + g_+(q^2) \frac{M_{\Lambda_b^0} - m_F}{s_-} \left( p_\mu + p'_\mu - \frac{q_\mu}{q^2} (M_{\Lambda_b^0}^2 - m_F^2) \right) \\ & \left. + g_\perp(q^2) \left( \gamma_\mu + \frac{2m_F}{s_-} p_\mu - \frac{2M_{\Lambda_b^0}}{s_-} p'_\mu \right) \right\} u_{\Lambda_b^0}(p, s), \quad (39) \end{aligned}$$

with the kinematic constraint  $g_0(0) = g_+(0)$  at  $q^2 \rightarrow 0$ .

The form factors have been discussed in the heavy quark limit; assuming both  $b$  and  $c$  as heavy, all the form factors  $f_i$  and  $g_i$  turn out to be identical [145]

$$f_0 = f_+ = f_\perp = g_0 = g_+ = g_\perp = \xi_B \quad (40)$$

and equal to the Isgur Wise function  $\xi_B$  for baryons. In the limit of a light baryon in the final state, the number of independent form factors is still reduced to two through the heavy quark symmetries of the  $\Lambda_b^0$ . It should be noted that the  $\Lambda_b^0 \rightarrow (p/\Lambda_c^+) \mu \nu$  decay rates peak at high  $q^2$ , which facilitates both lattice QCD calculations and experimental measurements.

The form factors for  $\Lambda_b^0$  decays have been studied on the lattice [146]. Based on these results the differential rates for both  $\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \nu$  as well as for  $\Lambda_b^0 \rightarrow p \mu \nu$  can be predicted in the full phase space. In particular, for the experimentally interesting region they find the ratio of decay rates to be [146]

$$\frac{\mathcal{B}(\Lambda_b^0 \rightarrow p \mu \bar{\nu})_{q^2 > 15 \text{ GeV}^2}}{\mathcal{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \bar{\nu})_{q^2 > 7 \text{ GeV}^2}} = (1.471 \pm 0.095 \pm 0.109) \left| \frac{V_{ub}}{V_{cb}} \right|^2 \quad (41)$$

where the first uncertainty is statistical and the second, systematic.

### Measurements at LHCb

The LHCb experiment has measured the branching fractions of the semileptonic decays  $\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \bar{\nu}$  and  $\Lambda_b^0 \rightarrow p \mu \bar{\nu}$ , from which they determine  $|V_{ub}|/|V_{cb}|$ . This is the first such determination at a hadron collider, the first to use a  $b$  baryon decay, and the first observation of  $\Lambda_b^0 \rightarrow p \mu \bar{\nu}$ . Excellent vertex resolution allows the  $p\mu$  and production vertices to be separated, which permits the calculation of the transverse momentum  $p_\perp$  of the  $p\mu$  pair relative to the  $\Lambda_b^0$  flight direction. The corrected mass,  $m_{\text{corr}} = \sqrt{p_\perp^2 + m_{p\mu}^2}$ , peaks at the  $\Lambda_b^0$  mass for signal decays and provides good discrimination against background combinations. The topologically similar decay  $\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \bar{\nu}$  is also measured, which eliminates the need to know the production cross-section or absolute efficiencies. Using vertex and  $\Lambda_b^0$  mass constraints,  $q^2$  can be determined up to a two-fold ambiguity. The LHCb analysis requires both solutions to be in

the high  $q^2$  region to minimise contamination from the low  $q^2$  region. They find [147]

$$\frac{\mathcal{B}(\Lambda_b^0 \rightarrow p \mu \bar{\nu})_{q^2 > 15 \text{ GeV}^2}}{\mathcal{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \bar{\nu})_{q^2 > 7 \text{ GeV}^2}} = (1.00 \pm 0.04 \pm 0.08) \times 10^{-2} \quad (42)$$

The largest systematic uncertainty is from knowledge of the branching fraction  $\mathcal{B}(\Lambda_c^+ \rightarrow p K^- \pi^+)$ ; uncertainties due to trigger, tracking and the  $\Lambda_c^+$  selection efficiency are each about 3%.

### The ratio $|V_{ub}|/|V_{cb}|$

The ratio of matrix elements,  $|V_{ub}|/|V_{cb}|$ , is often required when testing the compatibility of a set of measurements with theoretical predictions. It can be determined from the ratio of branching fractions measured by the LHCb experiment, quoted in the previous section. It can also be calculated based on the  $|V_{ub}|$  and  $|V_{cb}|$  values quoted earlier in this review.

As previously noted, the decay rate for  $\Lambda_b^0 \rightarrow p \mu \bar{\nu}$  peaks at high  $q^2$  where the calculation of the associated form factors using lattice QCD is under good control. Using the measured ratio from Eq. (42) along with the calculations of Ref. 146 results in

$$|V_{ub}|/|V_{cb}| = 0.083 \pm 0.006 \quad (\text{LHCb}). \quad (43)$$

Given the similarities in the theoretical frameworks used for charmed and charmless decays, we choose to quote the ratio  $|V_{ub}|/|V_{cb}|$  separately for inclusive and exclusive decays:

$$|V_{ub}|/|V_{cb}| = 0.107 \pm 0.006 \quad (\text{inclusive}), \quad (44)$$

$$|V_{ub}|/|V_{cb}| = 0.095 \pm 0.005 \quad (\text{exclusive}). \quad (45)$$

We average these values, along with the result in Eq. (43), weighting by relative errors. The average has  $p(\chi^2) = 1.8\%$ , so we scale the uncertainty by a factor 2.0 to find

$$|V_{ub}|/|V_{cb}| = 0.096 \pm 0.007 \quad (\text{average}). \quad (46)$$

### SEMITAUONIC DECAYS

*Summary:* Semileptonic decays to third-generation leptons provide sensitivity to non-Standard Model amplitudes, such as from a charged Higgs boson [148]. The ratios of branching fractions of semileptonic decays involving tau leptons to those involving  $e/\mu$ ,  $\mathcal{R}_{D^{(*)}} \equiv \mathcal{B}(\bar{B} \rightarrow D^{(*)} \tau \bar{\nu}_\tau) / \mathcal{B}(\bar{B} \rightarrow D^{(*)} \ell \bar{\nu}_\ell)$ , are predicted with good precision in the Standard Model [32,33,149,150]. We use the most precise values from lattice QCD for  $\mathcal{R}_D$  [32], and for  $\mathcal{R}_{D^*}$  we use a calculation based on the heavy quark expansion, combined with the measurements for  $\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell$  [150]

$$\begin{aligned} \mathcal{R}_D^{\text{SM}} &= 0.300 \pm 0.008, \\ \mathcal{R}_{D^*}^{\text{SM}} &= 0.252 \pm 0.003. \end{aligned} \quad (47)$$

Measurements [151–157] of these ratios yield higher values; averaging B-tagged measurements of  $\mathcal{R}_D$  and  $\mathcal{R}_{D^*}$  at the  $\Upsilon(4S)$  and the LHCb measurement of  $\mathcal{R}_D$  yields

$$\begin{aligned} \mathcal{R}_D^{\text{meas}} &= 0.391 \pm 0.050 \\ \mathcal{R}_{D^*}^{\text{meas}} &= 0.322 \pm 0.022 \end{aligned} \quad (48)$$

These values exceed Standard Model predictions by  $1.7\sigma$  and  $3.0\sigma$ , respectively. A variety of new physics models have been proposed [148,156–160] to explain this excess. The potential impact of any new physics in this decay mode on the  $|V_{ub}|$  and  $|V_{cb}|$  results given above is expected to be negligible.

#### Sensitivity of $\bar{B} \rightarrow D^{(*)}\tau\bar{\nu}_\tau$ to additional amplitudes

In addition to the helicity amplitudes present for decays to  $e\bar{\nu}_e$  and  $\mu\bar{\nu}_\mu$ , decays proceeding through  $\tau\bar{\nu}_\tau$  include a scalar amplitude  $H_s$ . The differential decay rate is given by [161]

$$\frac{d\Gamma}{dq^2} = \frac{G_F^2 |V_{cb}|^2 |\mathbf{p}_{D^{(*)}}^*|^2}{96\pi^3 m_B^2} \left(1 - \frac{m_\tau^2}{q^2}\right)^2 \left[ (|H_+|^2 + |H_-|^2 + |H_0|^2) \left(1 + \frac{m_\tau^2}{2q^2}\right) + \frac{3m_\tau^2}{2q^2} |H_s|^2 \right], \quad (49)$$

where  $|\mathbf{p}_{D^{(*)}}^*|$  is the 3-momentum of the  $D^{(*)}$  in the  $\bar{B}$  rest frame and the helicity amplitudes  $H$  depend on the four-momentum transfer  $q^2$ . All four helicity amplitudes contribute to  $\bar{B} \rightarrow D^*\tau\bar{\nu}_\tau$ , while only  $H_0$  and  $H_s$  contribute to  $\bar{B} \rightarrow D\tau\bar{\nu}_\tau$ ; as a result, new physics contributions tend to produce larger effects in the latter mode.

The (semi)-leptonic  $B$  decays into a  $\tau$  lepton provide a stringent test of the two-Higgs doublet model of type II (2HDMII), i.e. where the two Higgs doublets couple separately to up- and down-type quarks. This is also of relevance for Supersymmetry, since this corresponds to the Higgs sector of any commonly used supersymmetric model. These models involve additional charged scalar particles, which contribute at tree level to the (semi)-leptonic  $B$  decays into a  $\tau$ . The distinct feature of the 2HDMII is that the contributions of the charged scalars scale as  $m_\tau^2/m_{H^\pm}^2$ , since the couplings to the charged Higgs particles are proportional to the mass of the lepton. As a consequence, one may expect visible effects in decays into a  $\tau$ , but only small effects for decays into  $e$  and  $\mu$ .

As discussed in the next section, the 2HDMII does not describe the observations any better than the Standard Model. To achieve a better description one has to extend the analysis to other models, where the scaling of the new contributions with the lepton mass is different.

#### Measurement of $\mathcal{R}_{D^{(*)}}$

$\bar{B} \rightarrow D^{(*)}\tau\bar{\nu}_\tau$  decays have been studied at the  $\Upsilon(4S)$  resonance and in  $pp$  collisions.

At the  $\Upsilon(4S)$ , the experimental signature consists of a  $D$  or  $D^*$  meson, an electron or muon (denoted here by  $\ell$ ) from the decay  $\tau \rightarrow \ell\nu_\tau\bar{\nu}_\ell$ , a fully-reconstructed hadronic decay of the second  $B$  meson in the event, and multiple missing neutrinos. The signal decays are separated from  $\bar{B} \rightarrow D^{(*)}\ell\bar{\nu}_\ell$  decays using the lepton momentum and the measured missing mass squared; decays with only a single missing neutrino peak sharply at zero in this variable, while the signal is spread out to positive values. Background from  $\bar{B} \rightarrow D^{**}\ell\bar{\nu}_\ell$  decays with one or more unreconstructed particles is harder to separate from signal, as is background from  $\bar{B} \rightarrow D^{(*)}H_c X$  (where  $H_c$  is a hadron containing a  $\bar{c}$  quark) decays. The leading sources of

systematic uncertainty are due to the limited size of simulation samples used in constructing the PDFs, the composition of the  $D^{**}$  states, efficiency corrections, and cross-feed (swapping soft particles between the signal and tag  $B$ ).

The LHCb experiment has studied the decay  $\bar{B} \rightarrow D^{*+}\tau\bar{\nu}_\tau$  with  $D^{*+} \rightarrow D^0\pi^+$ ,  $D^0 \rightarrow K^-\pi^+$  and  $\tau \rightarrow \mu\nu_\tau\bar{\nu}_\mu$  in  $pp$  collisions. Their analysis takes advantage of the measurable flight lengths of  $b$  and  $c$  hadrons and  $\tau$  leptons. A multivariate discriminant is used to select decays where no additional charged particles are consistent with coming from the signal decay vertices. The separation between the primary and  $B$  decay vertices is used to calculate the momentum of the  $B$  decay products transverse to the  $B$  flight direction. The longitudinal component of the  $B$  momentum can be estimated based on the visible decay products; this allows a determination of the  $B$  rest frame, with modest resolution, and enables the calculation of the same discrimination variables available at the  $e^+e^- B$  factories. The (rest frame) muon energy, missing mass-squared and  $q^2$  are used in a 3- $d$  fit. The leading sources of systematic uncertainty are due to the size of the simulation sample used in constructing the fit templates, the shape of the muon misidentification template, and uncertainties in modelling the background from  $\bar{B} \rightarrow D^{**}\ell\bar{\nu}_\ell$  and  $\bar{B} \rightarrow D^{(*)}H_c X$  decays.

Measurements from Belle [152–154], BABAR [155,151] and LHCb [157] result in values for  $\mathcal{R}_D$  and  $\mathcal{R}_{D^*}$  that exceed Standard Model predictions. Table 3 lists these values and their average. The measurements of  $\mathcal{R}_D$  and  $\mathcal{R}_{D^*}$  have linear correlation coefficients of  $-0.27$  (BABAR) and  $-0.49$  (Belle); the averaged values have a correlation of  $-0.29$ . Two untagged Belle measurements [152,153] are subject to larger systematic uncertainties; we do not include them in our average.

**Table 3:** Measurements of  $\mathcal{R}_D$  and  $\mathcal{R}_{D^*}$ . The averages correspond to the values in the upper portion of the table.

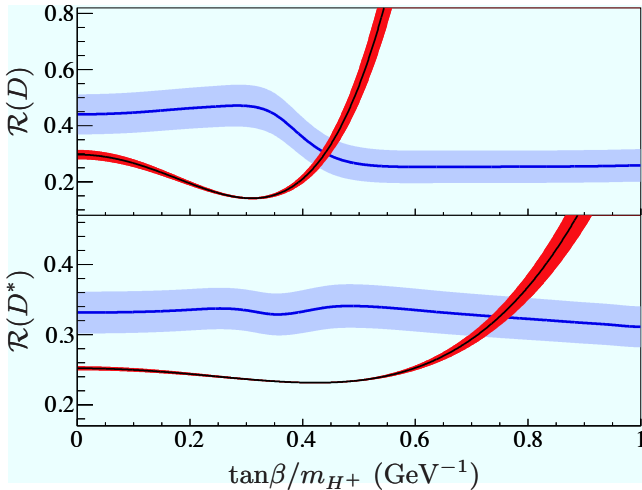
	$\mathcal{R}_D \times 10^2$	$\mathcal{R}_{D^*} \times 10^2$
BABAR [151] $B^0, B^+$	$44.0 \pm 5.8 \pm 4.2$	$33.2 \pm 2.4 \pm 1.8$
Belle [154] $B^0, B^+$	$37.5 \pm 6.4 \pm 2.6$	$29.3 \pm 3.8 \pm 1.5$
LHCb [157] $B^0$		$33.6 \pm 2.7 \pm 3.0$
Average $B^0, B^+$	$39.1 \pm 4.1 \pm 2.8$	$32.2 \pm 1.8 \pm 1.2$

The tension between the SM prediction and the measurements is at the level of  $1.7\sigma$  ( $\mathcal{R}_D$ ) and  $3.0\sigma$  ( $\mathcal{R}_{D^*}$ ); if one considers these deviations together the significance rises to  $3.9\sigma$ . This motivates speculation on possible new physics contributions. It is striking that an interpretation in terms of the 2HDMII seems to be ruled out by the data. Fig. 2 shows that the interpretation of the deviation of  $\mathcal{R}_D$  in terms of the 2HDMII requires vastly different values of the relevant parameter  $\tan\beta/m_{H^\pm}$  than for  $\mathcal{R}_{D^*}$ , excluding this possibility. All three experiments assume the Standard Model kinematic distributions for  $\bar{B} \rightarrow D^{(*)}\tau\bar{\nu}_\tau$  in their determinations of the branching fraction ratio. The BABAR [151] and Belle [154] analyses use

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## $V_{cb}$ and $V_{ub}$ CKM Matrix Elements

kinematical distributions from the 2HDMII when comparing the compatibility of their measurements with predictions; this is why the band in Fig. 2 corresponding to the measurement varies with  $\tan\beta/m_{H^+}$ . In general, new physics contributions with a different operator structure to the SM could modify  $R_{D^{(*)}}$  from the measured values, and could have a different effect in different experiments.



**Figure 2:** The  $R_{D^{(*)}}$  measured in Ref. 151 along with expectations in the 2HDMII as a function of  $\tan\beta/m_{H^+}$ .

A more general approach has been formulated in Ref. 159 on the basis of an effective field theory. Assuming lepton-flavour-universality-violating operators of dimension six and eight, the coefficients of these operators can be fitted to the observed values. Although a detailed analysis along these lines requires more data on related decays (such as  $B \rightarrow \pi\tau\bar{\nu}$ ), there are indications that the tension in  $R_{D^{(*)}}$  cannot be explained by a minimally flavor-violating scenario with only left-handed interactions; a better fit is obtained once right-handed and scalar currents are included.

## CONCLUSION

The study of semileptonic  $B$  meson decays continues to be an active area for both theory and experiment. The application of HQE calculations to inclusive decays is mature, and fits to moments of  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  decays provide precise values for  $|V_{cb}|$  and, in conjunction with input on  $m_c$  or from  $B \rightarrow X_s \gamma$  decays, provide precise and consistent values for  $m_b$ .

The determination of  $|V_{ub}|$  from inclusive  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  decays is based on multiple calculational approaches and independent measurements over a variety of kinematic regions, all of which provide consistent results. Further progress in this area is possible, but will require better theoretical control over higher-order terms, improved experimental knowledge of the  $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$  background and improvements to the modeling of the  $\bar{B} \rightarrow X_u \ell \bar{\nu}_\ell$  signal distributions.

In both  $b \rightarrow u$  and  $b \rightarrow c$  exclusive channels there has been significant recent progress in lattice-QCD calculations, resulting in improved precision on both  $|V_{ub}|$  and  $|V_{cb}|$ . These calculations now provide information on the form factors well away from the high  $q^2$  region, allowing better use of experimental data. Projections for future uncertainties from lattice calculations can be found in Ref. 162.

The values from the inclusive and exclusive determinations of both  $|V_{cb}|$  and  $|V_{ub}|$  are only marginally consistent. This is a long-standing puzzle, and the new measurement of  $|V_{ub}|/|V_{cb}|$  from LHCb based on  $\Lambda_b^0$  decays does not simplify the picture.

Both  $|V_{cb}|$  and  $|V_{ub}|$  are indispensable inputs into unitarity triangle fits. In particular, knowing  $|V_{ub}|$  with good precision allows a test of CKM unitarity in a most direct way, by comparing the length of the  $|V_{ub}|$  side of the unitarity triangle with the measurement of  $\sin(2\beta)$ . This comparison of a “tree” process ( $b \rightarrow u$ ) with a “loop-induced” process ( $B^0 - \bar{B}^0$  mixing) provides sensitivity to possible contributions from new physics.

The observation of semileptonic decays into  $\tau$  leptons has opened a new window to the physics of the third generation. The measurements indicate a tension between the data and the Standard Model prediction, which could be a hint to new physics. However, the most prominent and simplest candidate, the 2HDMII, cannot explain the current data. More general ansatzes fit the data, but do not provide deeper insight until measurements of related processes (such as  $B \rightarrow \pi\tau\bar{\nu}$ ) are available.

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## References

1. See R. Kowalewski and T. Mannel in K.A. Olive *et al.* (Particle Data Group), *Chin. Phys. C* **38**, 090001 (2014).
2. See “Heavy-Quark and Soft-Collinear Effective Theory” by C.W. Bauer and M. Neubert in this *Review*.
3. See “Lattice Quantum Chromodynamics” by S. Hashimoto, J. Laiho, and S.R. Sharpe in this *Review*.
4. See “Production and Decay of b-Flavored Hadrons” by M. Kreps and Y. Kwon in this *Review*.
5. Y. Amhis *et al.* (HFAG), [arXiv:1412.7515](https://arxiv.org/abs/1412.7515).
6. [slac.stanford.edu/xorg/hfag/semi/pdg14/home.shtml](http://slac.stanford.edu/xorg/hfag/semi/pdg14/home.shtml).
7. N. Isgur and M.B. Wise, *Phys. Lett.* **B232**, 113 (1989); *ibid.* **B237**, 527 (1990).
8. M.A. Shifman and M.B. Voloshin, *Sov. J. Nucl. Phys.* **47**, 511 (1988) [*Yad. Fiz.* **47**, 801 (1988)].
9. M.E. Luke, *Phys. Lett.* **B252**, 447 (1990).
10. A.V. Manohar and M.B. Wise, *Camb. Monogr. Part. Phys. Nucl. Phys. Cosmol.* **10**, 1 (2000); H. Georgi, *Phys. Lett.* **B240**, 447 (1990); A.F. Falk *et al.*, *Nucl. Phys.* **B343**, 1 (1990); E. Eichten and B. Hill, *Phys. Lett.* **B234**, 511 (1990).
11. A. Sirlin, *Nucl. Phys.* **B196**, 83 (1982).
12. J.A. Bailey *et al.* [Fermilab Lattice and MILC Collab.], *Phys. Rev.* **D89**, 114504 (2014).

13. C.G. Boyd, B. Grinstein, and R.F. Lebed, Phys. Rev. **D56**, 6895 (1997); *ibid.*, Phys. Rev. Lett. **74**, 4603 (1995); C.G. Boyd and M.J. Savage, Phys. Rev. **D56**, 303 (1997).
14. I. Caprini *et al.*, Nucl. Phys. **B530**, 153 (1998).
15. A. Czarnecki and K. Melnikov, Nucl. Phys. **B505**, 65 (1997).
16. I.I.Y. Bigi *et al.*, Phys. Rev. **D52**, 196 (1995).
17. A. Kapustin *et al.*, Phys. Lett. **B375**, 327 (1996).
18. P. Gambino, T. Mannel, and N. Uraltsev, Phys. Rev. **D81**, 113002 (2010).
19. P. Gambino, T. Mannel, and N. Uraltsev, JHEP **1210**, 169 (2012).
20. D. Buskulic *et al.* (ALEPH Collab.), Phys. Lett. **B395**, 373 (1997).
21. G. Abbiendi *et al.* (OPAL Collab.), Phys. Lett. **B482**, 15 (2000).
22. P. Abreu *et al.* (DELPHI Collab.), Phys. Lett. **B510**, 55 (2001).
23. J. Abdallah *et al.* (DELPHI Collab.), Eur. Phys. J. **C33**, 213 (2004).
24. N.E. Adam *et al.* (CLEO Collab.), Phys. Rev. **D67**, 032001 (2003).
25. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D77**, 032002 (2008).
26. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **100**, 231803 (2008).
27. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D79**, 012002 (2009).
28. W. Dungen *et al.* (Belle Collab.), Phys. Rev. **D82**, 112007 (2010).
29. C.G. Boyd, B. Grinstein, and R.F. Lebed, Phys. Rev. Lett. **74**, 4603 (1995).
30. C. Bourrely, L. Lellouch, and I. Caprini, Phys. Rev. **D79**, 013008 (2012).
31. M. Atoui *et al.*, Eur. Phys. J. **C74**, 2861 (2014).
32. H. Na *et al.*, Phys. Rev. **D92**, 054510 (2015).
33. J. Bailey *et al.* [Fermilab Lattice and MILC Collab.], Phys. Rev. **D92**, 034506 (2015).
34. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **104**, 011802 (2010).
35. R. Glattauer *et al.* (Belle Collab.), [arXiv:1510.03657](https://arxiv.org/abs/1510.03657).
36. A.V. Manohar and M.B. Wise, Phys. Rev. **D49**, 1310 (1994).
37. I.I.Y. Bigi *et al.*, Phys. Rev. Lett. **71**, 496 (1993); Phys. Lett. **B323**, 408 (1994).
38. D. Benson *et al.*, Nucl. Phys. **B665**, 367 (2003).
39. M. Gremm and A. Kapustin, Phys. Rev. **D55**, 6924 (1997).
40. B.M. Dassinger, T. Mannel, and S. Turczyk, JHEP **0703**, 087 (2007).
41. I.I. Bigi, N. Uraltsev, and R. Zwicky, Eur. Phys. J. **C50**, 539 (2007).
42. T. Mannel, S. Turczyk, and N. Uraltsev, JHEP **1011**, 109 (2010).
43. A. Pak and A. Czarnecki, Phys. Rev. **D78**, 114015 (2008).
44. S. Biswas and K. Melnikov, JHEP **1002**, 089 (2010).
45. P. Gambino, JHEP **1109**, 055 (2011).
46. P. Gambino and N. Uraltsev, Eur. Phys. J. **C34**, 181 (2004).
47. V. Aquila *et al.*, Nucl. Phys. **B719**, 77 (2005).
48. T. Becher, H. Boos, and E. Lunghi, JHEP **0712**, 062 (2007).
49. A. Alberti *et al.*, Nucl. Phys. **B870**, 16 (2013).
50. A. Alberti *et al.*, JHEP **1401**, 147 (2014).
51. T. Mannel, A.A. Pivovarov, and D. Rosenthal, Phys. Rev. **D92**, 054025 (2015).
52. C. Breidenbach *et al.*, Phys. Rev. **D78**, 014022 (2008).
53. I. Bigi *et al.*, JHEP **1004**, 073 (2010).
54. A.H. Hoang *et al.*, Phys. Rev. **D59**, 074017 (1999).
55. H. Leutwyler, Phys. Lett. **B98**, 447 (1981); M.B. Voloshin, Sov. J. Nucl. Phys. **36**, 143 (1982).
56. A.H. Hoang, Phys. Rev. **D61**, 034005 (2000).
57. S.E. Csorna *et al.* (CLEO Collab.), Phys. Rev. **D70**, 032002 (2004).
58. A.H. Mahmood *et al.* (CLEO Collab.), Phys. Rev. **D70**, 032003 (2004).
59. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D69**, 111103 (2004).
60. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D69**, 111104 (2004).
61. C. Schwanda *et al.* (Belle Collab.), Phys. Rev. **D75**, 032005 (2007).
62. P. Urquijo *et al.* (Belle Collab.), Phys. Rev. **D75**, 032001 (2007).
63. J. Abdallah *et al.* (DELPHI Collab.), Eur. Phys. J. **C45**, 35 (2006).
64. D. Acosta *et al.* (CDF Collab.), Phys. Rev. **D71**, 051103 (2005).
65. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D81**, 032003 (2010).
66. A. Limosani *et al.* [Belle Collab.], Phys. Rev. Lett. **103**, 241801 (2009).
67. C. Schwanda *et al.* (Belle Collab.), Phys. Rev. **D78**, 032016 (2008).
68. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D72**, 052004 (2005).
69. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **97**, 171803 (2006).
70. S. Chen *et al.* (CLEO Collab.), Phys. Rev. Lett. **87**, 251807 (2001).
71. M. Battaglia *et al.* Phys. Lett. **B556**, 41 (2003).
72. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **93**, 011803 (2004).
73. O. Buchmüller and H. Flächer, [hep-ph/0507253](https://arxiv.org/abs/hep-ph/0507253); updated in Ref. 6.
74. C.W. Bauer *et al.*, Phys. Rev. **D70**, 094017 (2004); updated in Ref. 5.
75. P. Gambino and C. Schwanda, Phys. Rev. **D89**, 014022 (2014).
76. A. Alberti *et al.*, Phys. Rev. Lett. **114**, 061802 (2015).
77. See section 5.4.2 of M. Antonelli *et al.*, Phys. Reports **494**, 197 (2010).
78. B. Dehnadi, *et al.*, JHEP **1309**, 103 (2013).
79. I. Allison *et al.* (HPQCD Collab.), Phys. Rev. **D78**, 054513 (2008).
80. K.G. Chetyrkin *et al.*, Phys. Rev. **D80**, 074010 (2009).

## Meson Particle Listings

 $V_{cb}$  and  $V_{ub}$  CKM Matrix Elements

81. T. Ewerth, P. Gambino, and S. Nandi, Nucl. Phys. **B830**, 278 (2010).
82. A.H. Hoang *et al.*, Phys. Rev. **D59**, 074017 (1999).
83. N. Uraltsev, Int. J. Mod. Phys. **A14**, 4641 (1999).
84. M. Neubert, Phys. Rev. **D49**, 4623 (1994); *ibid.* **D49**, 3392 (1994).
85. I. Bigi *et al.*, Int. J. Mod. Phys. **A9**, 2467 (1994).
86. C. W. Bauer *et al.*, Phys. Rev. **D68**, 094001 (2003).
87. M. Neubert, Phys. Lett. **B513**, 88 (2001); Phys. Lett. **B543**, 269 (2002).
88. A.K. Leibovich *et al.*, Phys. Rev. **D61**, 053006 (2000); Phys. Rev. **D62**, 014010 (2000); Phys. Lett. **B486**, 86 (2000); Phys. Lett. **B513**, 83 (2001).
89. A.H. Hoang *et al.*, Phys. Rev. **D71**, 093007 (2005).
90. B. Lange *et al.*, JHEP **0510**, 084 (2005); B. Lange, JHEP **0601**, 104 (2006).
91. M. Neubert, Phys. Lett. **B612**, 13 (2005).
92. Z. Ligeti, I.W. Stewart, and F.J. Tackmann, Phys. Rev. **D78**, 114014 (2008).
93. B.O. Lange, M. Neubert, and G. Paz, Phys. Rev. **D72**, 073006 (2005).
94. P. Gambino *et al.*, JHEP **0710**, 058 (2007).
95. J.R. Andersen and E. Gardi, JHEP **0601**, 097 (2006).
96. C.W. Bauer, Z. Ligeti, and M. E. Luke, Phys. Rev. **D64**, 113004 (2001); Phys. Lett. **B479**, 395 (2000).
97. C. Greub, M. Neubert, and B.D. Pecjak, Eur. Phys. J. **C65**, 501 (2010).
98. M. Brucherseifer, F. Caola, and K. Melnikov, Phys. Lett. **B721**, 107 (2013).
99. T. Mannel and S. Recksiegel, Phys. Rev. **D60**, 114040 (1999).
100. I.I.Y. Bigi and N.G. Uraltsev, Nucl. Phys. **B423**, 33 (1994).
101. M.B. Voloshin, Phys. Lett. **B515**, 74 (2001).
102. Z. Ligeti, M. Luke, and A.V. Manohar, Phys. Rev. **D82**, 033003 (2010).
103. P. Gambino and J.F. Kamenik, Nucl. Phys. **B840**, 424 (2010).
104. R. Barate *et al.* (ALEPH Collab.), Eur. Phys. J. **C6**, 555 (1999).
105. M. Acciarri *et al.* (L3 Collab.), Phys. Lett. **B436**, 174 (1998).
106. G. Abbiendi *et al.* (OPAL Collab.), Eur. Phys. J. **C21**, 399 (2001).
107. P. Abreu *et al.* (DELPHI Collab.), Phys. Lett. **B478**, 14 (2000).
108. A. Bornheim *et al.* (CLEO Collab.), Phys. Rev. Lett. **88**, 231803 (2002).
109. A. Limosani *et al.* (Belle Collab.), Phys. Lett. **B621**, 28 (2005).
110. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D73**, 012006 (2006).
111. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **95**, 111801 (2005), Erratum: Phys. Rev. Lett. **97**, 019903 (2006).
112. R. Kowalewski and S. Menke, Phys. Lett. **B541**, 29 (2002).
113. J.P. Lees *et al.* (BABAR Collab.), arXiv:1112.0702.
114. I. Bizjak *et al.* (Belle Collab.), Phys. Rev. Lett. **95**, 241801 (2005).
115. P. Urquijo *et al.* (Belle Collab.), Phys. Rev. Lett. **104**, 021801 (2010).
116. V. Golubev, Y. Skovpen, and V. Luth, Phys. Rev. **D76**, 114003 (2007).
117. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **96**, 221801 (2006).
118. H. Kakuno *et al.* (Belle Collab.), Phys. Rev. Lett. **92**, 101801 (2004).
119. A. Sibidanov *et al.* (Belle Collab.), Phys. Rev. **D88**, 032005 (2013).
120. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **90**, 181801 (2003).
121. T. Hokuue *et al.* (Belle Collab.), Phys. Lett. **B648**, 139 (2007).
122. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D79**, 052011 (2008).
123. C. Schwanda *et al.* (Belle Collab.), Phys. Rev. Lett. **93**, 131803 (2004).
124. N. E. Adam *et al.* (CLEO Collab.), Phys. Rev. Lett. **99**, 041802 (2007); Phys. Rev. **D76**, 012007 (2007); supercedes Phys. Rev. **D68**, 072003 (2003).
125. P. del Amo Sanchez *et al.*, (BABAR Collab.), Phys. Rev. **D83**, 032007 (2011); supercedes B. Aubert *et al.* (BABAR Collab.), Phys. Rev. **D72**, 051102 (2005).
126. P. del Amo Sanchez *et al.*, (BABAR Collab.), Phys. Rev. **D83**, 052011 (2011); updated in J.P. Lees *et al.* (BABAR Collab.), Phys. Rev. **D86**, 092004 (2012).
127. J.A. Bailey *et al.* (Fermilab/MILC Collab.), Phys. Rev. **D92**, 014024 (2015).
128. C.M. Bouchard *et al.*, Phys. Rev. **D90**, 054506 (2014).
129. J.M. Flynn *et al.*, Phys. Rev. **D91**, 074510 (2015).
130. T. Becher and R. J. Hill, Phys. Lett. **B633**, 61 (2006).
131. M.C. Arnesen *et al.*, Phys. Rev. Lett. **95**, 071802 (2005).
132. M.A. Shifman, A.I. Vainshtein, and V.I. Zakharov, Nucl. Phys. **B147**, 385 (1979); Nucl. Phys. **B147**, 448 (1979).
133. P. Ball and R. Zwicky, Phys. Rev. **D71**, 014015 (2005).
134. G. Duplancic *et al.*, JHEP **0804**, 014 (2008).
135. A. Khodjamirian *et al.*, Phys. Rev. **D83**, 094031 (2011).
136. A. Bharucha, JHEP **1205**, 092 (2012).
137. I.S. Imson *et al.*, JHEP **1502**, 126 (2015).
138. C. Bourrely, I. Caprini, and L. Lellouch, Phys. Rev. **D79**, 013008 (2009).
139. H. Ha *et al.* (Belle Collab.), Phys. Rev. **D83**, 071101 (2011).
140. K. Abe *et al.* (Belle Collab.), Phys. Lett. **B648**, 139 (2007).
141. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **101**, 081801 (2008).
142. B. Aubert *et al.* (BABAR Collab.), Phys. Rev. Lett. **97**, 211801 (2006).
143. W. Brower and H. Paar, Nucl. Instrum. Methods **A421**, 411 (1999).
144. P. Ball, arXiv:0705.2290; J.M. Flynn and J. Nieves, Phys. Lett. **B649**, 269 (2007); T. Becher and R.J. Hill, Phys. Lett. **B633**, 61 (2006); M. Arnesen *et al.*, Phys. Rev. Lett. **95**, 071802 (2005).

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# Meson Particle Listings

## $V_{cb}$ and $V_{ub}$ CKM Matrix Elements

145. T. Feldmann and M.W.Y. Yip, Phys. Rev. Lett. **85**, 014035 (2012).
146. W. Detmold, C. Lehner, and S. Meinel, Phys. Rev. **D92**, 034503 (2015).
147. The LHCb Collab., Nature Physics **11**, 743 (2015).
148. M. Tanaka, Z. Phys. **C67**, 321 (1995);  
H. Itoh, S. Komine, and Y. Okada, Prog. Theor. Phys. **114**, 179 (2005);  
U. Nierste, S. Trine, and S. Westhoff, Phys. Rev. **D78**, 015006 (2008);  
M. Tanaka and R. Watanabe, Phys. Rev. **D82**, 034027 (2010).
149. J.F. Kamenik and F. Mescia, Phys. Rev. **D78**, 014003 (2008).
150. S. Fajfer, J. F. Kamenik, and I. Nišandžić, Phys. Rev. **D85**, 094025 (2012).
151. J. Lees *et al.* (Babar Collab.), Phys. Rev. Lett. **109**, 101802 (2012); Phys. Rev. **D88**, 072012 (2013).
152. A. Bozek *et al.* (Belle Collab.), Phys. Rev. **D82**, 072005 (2010).
153. A. Matyja *et al.* (Belle Collab.), Phys. Rev. Lett. **99**, 191807 (2007).
154. M. Huschle *et al.* (Belle Collab.), Phys. Rev. **D92**, 072014 (2015).
155. B. Aubert *et al.* (Babar Collab.), Phys. Rev. Lett. **100**, 021801 (2008).
156. A. Datta, M. Duraisamy, and D. Ghosh, Phys. Rev. **D86**, 034027 (2012).
157. The LHCb Collab., Phys. Rev. Lett. **115**, 111803 (2015).
158. D. Becirevic, N. Kosnik, and A. Tayduganov, Phys. Lett. **B716**, 208 (2012).
159. S. Fajfer *et al.*, Phys. Rev. Lett. **109**, 161801 (2012).
160. A. Crivellin, C. Greub, and A. Kokulu, Phys. Rev. **D86**, 054014 (2012).
161. J. G. Körner and G. A. Schuler, Z. Phys. **C46**, 93 (1990).
162. USQCD Collab. (2011),  
[www.usqcd.org/documents/HiIntensityFlavor.pdf](http://www.usqcd.org/documents/HiIntensityFlavor.pdf).

0.0431 ± 0.0013 ± 0.0018	<sup>5</sup> ADAM	03	CLE2	$e^+e^- \rightarrow \gamma(4S)$
0.0355 ± 0.0014 ± $\begin{smallmatrix} +0.0023 \\ -0.0024 \end{smallmatrix}$	<sup>6</sup> ABREU	01H	DLPH	$e^+e^- \rightarrow Z$
0.0371 ± 0.0010 ± 0.0020	<sup>7</sup> ABBIENDI	00Q	OPAL	$e^+e^- \rightarrow Z$
0.0319 ± 0.0018 ± 0.0019	<sup>8</sup> BUSKULIC	97	ALEP	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0344 ± 0.0003 ± 0.0011	<sup>9</sup> AUBERT	08R	BABR	Repl. by AUBERT 09A
0.0355 ± 0.0003 ± 0.0016	<sup>10</sup> AUBERT	05E	BABR	Repl. by AUBERT 08R
0.0377 ± 0.0011 ± 0.0019	<sup>11</sup> ABDALLAH	04D	DLPH	$e^+e^- \rightarrow Z^0$
0.0354 ± 0.0019 ± 0.0018	<sup>12</sup> ABE	02F	BELL	Repl. by DUNGEL 10
0.0431 ± 0.0013 ± 0.0018	<sup>13</sup> BRIERE	02	CLE2	$e^+e^- \rightarrow \gamma(4S)$
0.0328 ± 0.0019 ± 0.0022		97G	OPAL	Repl. by ABBIENDI 00Q
0.0350 ± 0.0019 ± 0.0023	<sup>14</sup> ABREU	96P	DLPH	Repl. by ABREU 01H
0.0351 ± 0.0019 ± 0.0020	<sup>15</sup> BARISH	95	CLE2	Repl. by ADAM 03
0.0314 ± 0.0023 ± 0.0025	BUSKULIC	95N	ALEP	Repl. by BUSKULIC 97

<sup>1</sup> Uses fully reconstructed  $D^{*-}\ell^+\nu$  events ( $\ell = e$  or  $\mu$ ).

<sup>2</sup> Obtained from a global fit to  $B \rightarrow D^{(*)}\ell\nu_\ell$  events, with reconstructed  $D^0\ell$  and  $D^+\ell$  final states and  $\rho^2 = 1.22 \pm 0.02 \pm 0.07$ .

<sup>3</sup> Measured using the dependence of  $B^- \rightarrow D^{*0}e^-\bar{\nu}_e$  decay differential rate and the form factor description by CAPRINI 98 with  $\rho^2 = 1.16 \pm 0.06 \pm 0.08$ .

<sup>4</sup> Measurement using fully reconstructed  $D^*$  sample with a  $\rho^2 = 1.32 \pm 0.15 \pm 0.33$ .

<sup>5</sup> Average of the  $B^0 \rightarrow D^{*}(2010)^-\ell^+\nu$  and  $B^+ \rightarrow \bar{D}^{*}(2007)\ell^+\nu$  modes with  $\rho^2 = 1.61 \pm 0.09 \pm 0.21$  and  $f_{+,-} = 0.521 \pm 0.012$ .

<sup>6</sup> ABREU 01H measured using about 5000 partial reconstructed  $D^*$  sample with a  $\rho^2 = 1.34 \pm 0.14 \pm \begin{smallmatrix} +0.24 \\ -0.22 \end{smallmatrix}$ .

<sup>7</sup> ABBIENDI 00Q: measured using both inclusively and exclusively reconstructed  $D^{*\pm}$  samples with a  $\rho^2 = 1.21 \pm 0.12 \pm 0.20$ . The statistical and systematic correlations between  $|V_{cb}| \times F(1)$  and  $\rho^2$  are 0.90 and 0.54 respectively.

<sup>8</sup> BUSKULIC 97: measured using exclusively reconstructed  $D^{*\pm}$  with a  $a^2 = 0.31 \pm 0.17 \pm 0.08$ . The statistical correlation is 0.92.

<sup>9</sup> Measured using fully reconstructed  $D^*$  sample and a simultaneous fit to the Caprini-Lellouch-Neubert form factor parameters:  $\rho^2 = 1.191 \pm 0.048 \pm 0.028$ ,  $R_1(1) = 1.429 \pm 0.061 \pm 0.044$ , and  $R_2(1) = 0.827 \pm 0.038 \pm 0.022$ .

<sup>10</sup> Measurement using fully reconstructed  $D^*$  sample with a  $\rho^2 = 1.29 \pm 0.03 \pm 0.10$ .

<sup>11</sup> Combines with previous partial reconstructed  $D^*$  measurement with a  $\rho^2 = 1.39 \pm 0.10 \pm 0.33$ .

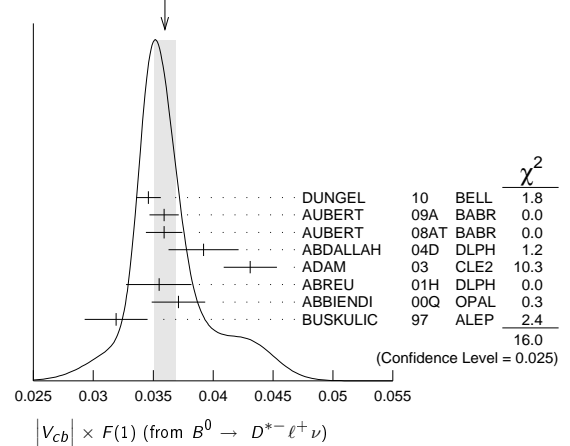
<sup>12</sup> Measured using exclusive  $B^0 \rightarrow D^{*}(892)^-\ell^+\nu$  decays with  $\rho^2 = 1.35 \pm 0.17 \pm 0.19$  and a correlation of 0.91.

<sup>13</sup> BRIERE 02 result is based on the same analysis and data sample reported in ADAM 03.

<sup>14</sup> ABREU 96P: measured using both inclusively and exclusively reconstructed  $D^{*\pm}$  samples.

<sup>15</sup> BARISH 95: measured using both exclusive reconstructed  $B^0 \rightarrow D^{*-}\ell^+\nu$  and  $B^+ \rightarrow D^{*0}\ell^+\nu$  samples. They report their experiment's uncertainties  $\pm 0.0019 \pm 0.0018 \pm 0.0008$ , where the first error is statistical, the second is systematic, and the third is the uncertainty in the lifetimes. We combine the last two in quadrature.

WEIGHTED AVERAGE  
0.0360 ± 0.0009 (Error scaled by 1.5)



### $V_{cb}$ MEASUREMENTS

For the discussion of  $V_{cb}$  measurements, which is not repeated here, see the review on "Determination of  $|V_{cb}|$  and  $|V_{ub}|$ ."

The CKM matrix element  $|V_{cb}|$  can be determined by studying the rate of the semileptonic decay  $B \rightarrow D^{(*)}\ell\nu$  as a function of the recoil kinematics of  $D^{(*)}$  mesons. Taking advantage of theoretical constraints on the normalization and a linear  $\omega$  dependence of the form factors ( $F(\omega)$ ,  $G(\omega)$ ) provided by Heavy Quark Effective Theory (HQET), the  $|V_{cb}| \times F(\omega)$  and  $\rho^2$  ( $a^2$ ) can be simultaneously extracted from data, where  $\omega$  is the scalar product of the two-meson four velocities,  $F(1)$  is the form factor at zero recoil ( $\omega=1$ ) and  $\rho^2$  is the slope, sometimes denoted as  $a^2$ . Using the theoretical input of  $F(1)$ , a value of  $|V_{cb}|$  can be obtained.

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

### $|V_{cb}| \times F(1)$ (from $B^0 \rightarrow D^{*-}\ell^+\nu$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.03581 ± 0.00045 OUR EVALUATION</b>	with $\rho^2 = 1.207 \pm 0.026$ and a correlation 0.324. The fitted $\chi^2$ is 30.0 for 23 degrees of freedom.		
<b>0.0360 ± 0.0009 OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.		
0.0346 ± 0.0002 ± 0.0010	<sup>1</sup> DUNGEL	10	BELL $e^+e^- \rightarrow \gamma(4S)$
0.0359 ± 0.0002 ± 0.0012	<sup>2</sup> AUBERT	09A	BABR $e^+e^- \rightarrow \gamma(4S)$
0.0359 ± 0.0006 ± 0.0014	<sup>3</sup> AUBERT	08AT	BABR $e^+e^- \rightarrow \gamma(4S)$
0.0392 ± 0.0018 ± 0.0023	<sup>4</sup> ABDALLAH	04D	DLPH $e^+e^- \rightarrow Z^0$

### $|V_{cb}| \times G(1)$ (from $B \rightarrow D^{*-}\ell^+\nu$ )

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.04265 ± 0.00153 OUR EVALUATION</b>	with $\rho^2 = 1.190 \pm 0.054$ and a correlation 0.83. The fitted $\chi^2$ is 0.5 for 8 degrees of freedom.		
<b>0.0422 ± 0.0010 OUR AVERAGE</b>			
0.04229 ± 0.00137	<sup>16</sup> GLATTAUER	16	BELL $e^+e^- \rightarrow \gamma(4S)$
0.0423 ± 0.0019 ± 0.0014	<sup>17</sup> AUBERT	10	BABR $e^+e^- \rightarrow \gamma(4S)$
0.0431 ± 0.0008 ± 0.0023	<sup>18</sup> AUBERT	09A	BABR $e^+e^- \rightarrow \gamma(4S)$
0.0416 ± 0.0047 ± 0.0037	<sup>19</sup> BARTELT	99	CLE2 $e^+e^- \rightarrow \gamma(4S)$
0.0278 ± 0.0068 ± 0.0065	<sup>20</sup> BUSKULIC	97	ALEP $e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.0411 ± 0.0044 ± 0.0052	<sup>21</sup> ABE	02E	BELL Repl. by GLATTAUER 16
0.0337 ± 0.0044 ± $\begin{smallmatrix} +0.0072 \\ -0.0049 \end{smallmatrix}$	<sup>22</sup> ATHANAS	97	CLE2 Repl. by BARTELT 99





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Meson Particle Listings

$B_1(5721)^0$ ,  $B^*_j(5732)$ ,  $B^*_2(5747)^+$

$B_1(5721)^0$

$I(J^P) = \frac{1}{2}(1^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

$B_1(5721)^0$  MASS

OUR FIT uses mass differences measurements listed below to determine the mass  $m_{B_1(5721)^0}$ .

VALUE (MeV)	DOCUMENT ID
<b>5726.0±1.3 OUR FIT</b>	Error includes scale factor of 1.2.

$m_{B_1^0} - m_{B^+}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>446.7±1.3 OUR FIT</b>	Error includes scale factor of 1.2.		
<b>441.5±2.4±1.3</b>	<sup>1</sup> ABAZOV	07T D0	$p\bar{p}$ at 1.96 TeV
• • •	We do not use the following data for averages, fits, limits, etc.	• • •	
446.2 <sup>+1.9+1.0</sup> <sub>-2.1-1.2</sub>	<sup>1</sup> AALTONEN	09D CDF	Repl. by AALTONEN 14i
<sup>1</sup> Observed in $B_1^0 \rightarrow B^{*+}\pi^-$ .			

$m_{B_1^0} - m_{B^{*+}}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>401.4±1.2 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>402.8±1.1 OUR AVERAGE</b>				
403.4±0.7±1.5	35K	<sup>2</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
402.3±0.9 <sup>+1.1</sup> <sub>-1.2</sub>		<sup>3</sup> AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV
<sup>2</sup> AAIJ 15AB reports $[m_{B_1^0} - m_{B^+}] - (m_{B^{*+}} - m_{B^+}) - m_{\pi^-} = 263.9 \pm 0.7 \pm 1.4$ MeV which we adjust by the $\pi^-$ mass and $(m_{B^{*+}} - m_{B^+}) = 45.01 \pm 0.30 \pm 0.23$ MeV. The masses inside the square brackets were measured for each candidate event.				
<sup>3</sup> AALTONEN 14i reports $m_{B_1(5721)^0} - m_{B^{*+}} - m_{\pi^-} = 262.7 \pm 0.9^{+1.1}_{-1.2}$ MeV which we adjusted by the $\pi^-$ mass.				

$B_1(5721)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>27.5±3.4 OUR AVERAGE</b>	Error includes scale factor of 1.1.			
30.1±1.5±3.5	35k	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
23 ±3 ±4		AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV

$B_1(5721)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $B^{*+}\pi^-$	dominant

$B_1(5721)^0$  BRANCHING RATIOS

$\Gamma(B^{*+}\pi^-)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
seen	35 K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV	
dominant		AALTONEN	09D CDF	$p\bar{p}$ at 1.96 TeV	
dominant		<sup>4</sup> ABAZOV	07T D0	$p\bar{p}$ at 1.96 TeV	
<sup>4</sup> Observed in $B_1^0 \rightarrow B^{*+}\pi^-$ with $B^{*+} \rightarrow B^+\gamma$ and $B^+ \rightarrow J/\psi\pi^+$ .					

$B_1(5721)^0$  REFERENCES

AAIJ	15AB JHEP 1504 024	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	14i PR D90 012013	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	09D PRL 102 102003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	07T PRL 99 172001	V.M. Abazov <i>et al.</i>	(D0 Collab.)

$B^*_j(5732)$   
or  $B^{**}$

$I(J^P) = ?(??)$   
 $I, J, P$  need confirmation.

OMITTED FROM SUMMARY TABLE  
Signal can be interpreted as stemming from several narrow and broad resonances. Needs confirmation.

$B^*_j(5732)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5698± 8 OUR AVERAGE</b>	Error includes scale factor of 1.2.			
5710±20		<sup>1</sup> AFFOLDER	01F CDF	$p\bar{p}$ at 1.8 TeV
5695 <sup>+17</sup> <sub>-19</sub>		<sup>2</sup> BARATE	98L ALEP	$e^+e^- \rightarrow Z$
5704± 4±10	1944	<sup>3</sup> BUSKULIC	96D ALEP	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
5732± 5±20	2157	ABREU	95B DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
5681±11	1738	AKERS	95E OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
5713± 2		<sup>4</sup> ACCIARRI	99N L3	$e^+e^- \rightarrow Z$

- <sup>1</sup> AFFOLDER 01F uses the reconstructed  $B$  meson through semileptonic decay channels. The fraction of light  $B$  mesons that are produced at  $L=1$   $B^{**}$  states is measured to be  $0.28 \pm 0.06 \pm 0.03$ .
- <sup>2</sup> BARATE 98L uses fully reconstructed  $B$  mesons to search for  $B^{**}$  production in the  $B\pi^\pm$  system. In the framework of heavy quark symmetry (HQS), they also measured the mass of  $B_2^*$  to be  $5739^{+8+6}_{-11-4}$  MeV/ $c^2$  and the relative production rate of  $B(b \rightarrow B_2^* \rightarrow B^{(*)}\pi)/B(b \rightarrow B_{u,d}) = (31 \pm 9^{+6}_{-5})\%$ .
- <sup>3</sup> Using  $m_{B\pi^-} - m_B = 424 \pm 4 \pm 10$  MeV.
- <sup>4</sup> ACCIARRI 99N uses inclusive reconstructed  $B$  mesons to search for  $B^{**}$  production in the  $B^{(*)}\pi^\pm$  system. In the framework of HQET, they measured the mass of  $B_1^*$  and  $B_2^*$  to be  $5670 \pm 10 \pm 13$  MeV and  $5768 \pm 5 \pm 6$  with the  $B(b \rightarrow B^{**}) = (32 \pm 3 \pm 6) \times 10^{-2}$ . They also reported the evidence for the existence of an excited  $B$ -meson state or mixture of states in the region 5.9–6.0 GeV.

$B^*_j(5732)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>128±18 OUR AVERAGE</b>				
145±28	2157	ABREU	95B DLPH	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV
116±24	1738	AKERS	95E OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

$B^*_j(5732)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $B^*\pi + B\pi$	dominant
$\Gamma_2$ $B^*\pi(X)$	[a] (85±29) %
[a] X refers to decay modes with or without additional accompanying decay particles.	

$B^*_j(5732)$  BRANCHING RATIOS

X refers to decay modes with or without additional accompanying decay particles.

$\Gamma(B^*\pi(X))/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$0.85^{+0.26}_{-0.27} \pm 0.12$	ABBIENDI	02E OPAL	$e^+e^- \rightarrow Z$

$B^*_j(5732)$  REFERENCES

ABBIENDI	02E EPJ C23 437	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
AFFOLDER	01F PR D64 072002	T. Affolder <i>et al.</i>	(CDF Collab.)
ACCIARRI	99N PL B465 323	M. Acciarri <i>et al.</i>	(L3 Collab.)
BARATE	98L PL B425 215	R. Barate <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	96D ZPHY C69 393	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ABREU	95B PL B345 598	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	95E ZPHY C66 19	R. Akers <i>et al.</i>	(OPAL Collab.)

$B^*_2(5747)^+$

$I(J^P) = \frac{1}{2}(2^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

$B^*_2(5747)^+$  MASS

OUR FIT uses  $m_{B^0}$  and  $m_{B^{*+}} - m_{B^0}$  to determine  $m_{B^*_2(5747)^+}$ .

VALUE (MeV)	DOCUMENT ID
<b>5737.2±0.7 OUR FIT</b>	

$m_{B_2^{*+}} - m_{B^0}$					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>457.5 ±0.7 OUR FIT</b>					
<b>457.5 ±0.7 OUR AVERAGE</b>					
457.62±0.72±0.40	4K	<sup>1</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV	
457.3 ±1.3 <sup>+0.3</sup> <sub>-0.9</sub>		<sup>2</sup> AALTONEN	14i CDF	$p\overline{p}$ at 1.96 TeV	

$B^*_2(5747)^+$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>20 ±5 OUR AVERAGE</b>	Error includes scale factor of 2.2.			
23.6±2.0±2.1	4K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
11 <sup>+4+3</sup> <sub>-3-4</sub>		AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV

Meson Particle Listings

$B_2^*(5747)^+, B_2^*(5747)^0, B_J(5840)^+$

$B_2^*(5747)^+$ DECAY MODES				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$B^0\pi^+$	seen		
$\Gamma_2$	$B^{*0}\pi^+$	seen		

$B_2^*(5747)^+$ BRANCHING RATIOS				
$\Gamma(B^0\pi^+)/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	4K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
seen		AALTONEN	14i CDF	$p\overline{p}$ at 1.96 TeV
$\Gamma(B^{*0}\pi^+)/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	4k	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
$\Gamma(B^{*0}\pi^+)/\Gamma(B^0\pi^+)$		$\Gamma_2/\Gamma_1$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$1.0\pm 0.5\pm 0.8$	4k	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV

$B_2^*(5747)^+$ REFERENCES				
AAIJ	15AB JHEP 1504 024	R. Aaij <i>et al.</i>	(LHCb Collab.)	
AALTONEN	14i PR D90 012013	T. Aaltonen <i>et al.</i>	(CDF Collab.)	

$B_2^*(5747)^0$

$I(J^P) = \frac{1}{2}(2^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

$B_2^*(5747)^0$ MASS				
OUR FIT uses $m_{B^+}, m_{B_1^0} - m_{B^+}$ , and $m_{B_2^{*0}} - m_{B_1^0}$ to determine $m_{B_2^*(5747)^0}$ . The $-0.659$ correlation between statistical uncertainties of $m_{B_1^0} - m_{B^+}$ and $m_{B_2^{*0}} - m_{B_1^0}$ measurements reported by ABAZOV 07T is taken into account.				
VALUE (MeV)	DOCUMENT ID			
<b>5739.5 <math>\pm</math> 0.7 OUR FIT</b>	Error includes scale factor of 1.4.			
$m_{B_2^{*0}} - m_{B_1^0}$				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>13.5 <math>\pm</math> 1.4 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>26.2 <math>\pm</math> 3.1 <math>\pm</math> 0.9</b>	<sup>1</sup> ABAZOV	07T D0	$p\overline{p}$ at 1.96 TeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
14.9 $^{+2.2}_{-2.5} \pm 1.2$	<sup>1</sup> AALTONEN	09D CDF	Repl. by AALTONEN 14i	
<sup>1</sup> Observed in $B_2^{*0} \rightarrow B^{*+} \pi^-$ and $B_2^{*0} \rightarrow B^+ \pi^-$ .				
$m_{B_2^{*0}} - m_{B^+}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>460.2 <math>\pm</math> 0.6 OUR FIT</b>	Error includes scale factor of 1.4.			
<b>459.9 <math>\pm</math> 0.8 OUR AVERAGE</b>	Error includes scale factor of 1.8.			
460.18 $\pm$ 0.37 $\pm$ 0.33	17K	<sup>2</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
457.5 $\pm$ 1.2 $^{+0.8}_{-0.9}$		<sup>3</sup> AALTONEN	14i CDF	$p\overline{p}$ at 1.96 TeV
<sup>2</sup> AAIJ 15AB reports $[m_{B_2^{*0}} - m_{B^+}] - m_{\pi^-} = 320.6 \pm 0.4 \pm 0.3$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event.				
<sup>3</sup> AALTONEN 14i reports $m_{B_2^*(5747)^0} - m_{B^+} - m_{\pi^-} = 317.9 \pm 1.2^{+0.8}_{-0.9}$ MeV which we adjusted by the $\pi^-$ mass.				

$B_2^*(5747)^0$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>24.2 <math>\pm</math> 1.7 OUR AVERAGE</b>				
24.5 $\pm$ 1.0 $\pm$ 1.5	17K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
22 $^{+3}_{-2} \pm 4$		AALTONEN	14i CDF	$p\overline{p}$ at 1.96 TeV
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
22.7 $^{+3.8}_{-3.2} \pm 3.2$		AALTONEN	09D CDF	Repl. by AALTONEN 14i

$B_2^*(5747)^0$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$	$B^+\pi^-$	
$\Gamma_2$	$B^{*+}\pi^-$	

$B_2^*(5747)^0$ BRANCHING RATIOS				
$\Gamma(B^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	17K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
dominant		AALTONEN	09D CDF	$p\overline{p}$ at 1.96 TeV
dominant		ABAZOV	07T D0	$p\overline{p}$ at 1.96 TeV
$\Gamma(B^{*+}\pi^-)/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	17K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
dominant		AALTONEN	09D CDF	$p\overline{p}$ at 1.96 TeV
dominant		ABAZOV	07T D0	$p\overline{p}$ at 1.96 TeV
$\Gamma(B^{*+}\pi^-)/\Gamma(B^+\pi^-)$		$\Gamma_2/\Gamma_1$		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.82 <math>\pm</math> 0.28 OUR AVERAGE</b>				
0.71 $\pm$ 0.14 $\pm$ 0.30	17K	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
1.10 $\pm$ 0.42 $\pm$ 0.31		<sup>4</sup> ABAZOV	07T D0	$p\overline{p}$ at 1.96 TeV
<sup>4</sup> Converted from measured ratio of $R = \text{B}(B_2^{*0} \rightarrow B^{*+}\pi^-) / \text{B}(B_2^{*0} \rightarrow B^{(*)+}\pi^-) = 0.475 \pm 0.095 \pm 0.069$ .				

$B_2^*(5747)^0$ REFERENCES				
AAIJ	15AB JHEP 1504 024	R. Aaij <i>et al.</i>	(LHCb Collab.)	
AALTONEN	14i PR D90 012013	T. Aaltonen <i>et al.</i>	(CDF Collab.)	
AALTONEN	09D PRL 102 102003	T. Aaltonen <i>et al.</i>	(CDF Collab.)	
ABAZOV	07T PRL 99 172001	V.M. Abazov <i>et al.</i>	(D0 Collab.)	

$B_J(5840)^+$

$I(J^P) = \frac{1}{2}(??)$  Status: \*\*  
 $I, J, P$  need confirmation.

OMITTED FROM SUMMARY TABLE  
Quantum numbers shown are quark-model predictions.

$B_J(5840)^+$ MASS				
OUR FIT uses $m_{B^0}$ and $m_{B_J(5840)^+} - m_{B^0}$ to determine $m_{B_J(5840)^+}$ .				
VALUE (MeV)	DOCUMENT ID			
<b>5851 ± 19 OUR FIT</b>				
$m_{B_J(5840)^+} - m_{B^0}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>571 ± 19 OUR FIT</b>				
<b>571 ± 13 ± 14</b>	7k	<sup>1</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
595 ± 26 ± 14	7k	<sup>2</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AB reports $[m_{B^{*+}} - m_{B^0}] - m_{\pi^+} = 431 \pm 13 \pm 14$ MeV which we adjust by the $\pi^+$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses two relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>2</sup> AAIJ 15AB reports $[m_{B^{*+}} - m_{B^0}] - m_{\pi^+} = 455 \pm 26 \pm 14$ MeV which we adjust by the $\pi^+$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses three relativistic Breit-Wigner functions in the fit for mass difference.				
$m_{B_J(5840)^+} - m_{B^{*0}}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
565 ± 15 ± 14	7k	<sup>3</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>3</sup> AAIJ 15AB reports $[m_{B^{*+}} - m_{B^0}] - (m_{B^{*+}} - m_{B^+}) - m_{\pi^+} = 425 \pm 15 \pm 14$ MeV which we adjust by the $\pi^+$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = -(-1)^J$ , $(m_{B^{*0}} - m_{B^0}) = (m_{B^{*+}} - m_{B^+}) = 45.01 \pm 0.30 \pm 0.23$ MeV, and uses three relativistic Breit-Wigner functions in the fit for mass difference.				

$B_J(5840)^+$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>224 <math>\pm</math> 24 <math>\pm</math> 80</b>	7k	<sup>4</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
• • •	We do not use the following data for averages, fits, limits, etc. • • •			
215 $\pm$ 27 $\pm$ 80	7k	<sup>5</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
229 $\pm$ 27 $\pm$ 80	7k	<sup>6</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>4</sup> Assuming $P = (-1)^J$ and using two relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>5</sup> Assuming $P = (-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>6</sup> Assuming $P = -(-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				

See key on page 601

Meson Particle Listings

$B_J(5840)^+$ ,  $B_J(5840)^0$ ,  $B_J(5970)^+$

$B_J(5840)^+$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $B^{*0}\pi^+$	seen	
$\Gamma_2$ $B^0\pi^+$	possibly seen	

$B_J(5840)^+$ BRANCHING RATIOS				
$\Gamma(B^{*0}\pi^+)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	7k	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
$\Gamma(B^0\pi^+)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
possibly seen	7k	<sup>7</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>7</sup> A $B\pi$ decay is forbidden from a $P = -(-1)^J$ parent, whereas $B^*\pi$ is allowed.				

$B_J(5840)^+$ REFERENCES			
AAIJ	15AB JHEP 1504 024	R. Aaij et al.	(LHCb Collab.)
<div><math>B_J(5840)^0</math><math>I(J^P) = \frac{1}{2}(?)^?</math> Status: ** <math>I, J, P</math> need confirmation.</div>			
OMITTED FROM SUMMARY TABLE			
Quantum numbers shown are quark-model predictions.			

$B_J(5840)^0$ MASS	
OUR FIT uses $m_{B^+}$ and $m_{B(5840)^0} - m_{B^+}$ to determine $m_{B_J(5840)^0}$ .	
VALUE (MeV)	DOCUMENT ID
<b>5863±9 OUR FIT</b>	

$m_{B_J(5840)^0} - m_{B^+}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>584±9 OUR FIT</b>				
<b>584±5±7</b>	12k	<sup>1</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
610±22±7	12k	<sup>2</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AB reports $[m_{B_J^0} - m_{B^+}] - m_{\pi^-} = 444 \pm 5 \pm 7$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses two relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>2</sup> AAIJ 15AB reports $[m_{B_J^0} - m_{B^+}] - m_{\pi^-} = 471 \pm 22 \pm 7$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses three relativistic Breit-Wigner functions in the fit for mass difference.				
$m_{B_J(5840)^0} - m_{B^{*+}}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
584±5±7	12k	<sup>3</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>3</sup> AAIJ 15AB reports $[m_{B_J^0} - m_{B^+}] - (m_{B^{*+}} - m_{B^+}) - m_{\pi^-} = 444 \pm 5 \pm 7$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = -(-1)^J$ , $(m_{B^{*+}} - m_{B^+}) = 45.01 \pm 0.30 \pm 0.23$ MeV, and uses three relativistic Breit-Wigner functions in the fit for mass difference.				

$B_J(5840)^0$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>127±17±34</b>	12k	<sup>4</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
107±20±34	12k	<sup>5</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
119±17±34	12k	<sup>6</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>4</sup> Assuming $P = (-1)^J$ and using two relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>5</sup> Assuming $P = (-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>6</sup> Assuming $P = -(-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				

$B_J(5840)^0$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $B^{*+}\pi^-$	seen	
$\Gamma_2$ $B^+\pi^-$	possibly seen	

$B_J(5840)^0$ BRANCHING RATIOS				
$\Gamma(B^{*+}\pi^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	12k	AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
$\Gamma(B^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
possibly seen	<sup>7</sup> AAIJ	15AB LHCb		$pp$ at 7, 8 TeV
<sup>7</sup> A $B\pi$ decay is forbidden from a $P = -(-1)^J$ parent, whereas $B^*\pi$ is allowed.				

$B_J(5840)^0$ REFERENCES			
AAIJ	15AB JHEP 1504 024	R. Aaij et al.	(LHCb Collab.)
<div><math>B_J(5970)^+</math><math>I(J^P) = \frac{1}{2}(?)^?</math> Status: ** <math>I, J, P</math> need confirmation.</div>			
Quantum numbers shown are quark-model predictions.			

$B_J(5970)^+$ MASS	
OUR FIT uses $m_{B^0}$ and $m_{B_J(5970)^+} - m_{B^0}$ to determine $m_{B_J(5970)^+}$ .	
VALUE (MeV)	DOCUMENT ID
<b>5964±5 OUR FIT</b>	

$m_{B_J(5970)^+} - m_{B^0}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>685±5 OUR FIT</b>				
<b>685±5 OUR AVERAGE</b>				
685.3±4.1±2.5	2K	<sup>1</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
681±5±12	1.4k	<sup>2</sup> AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
686.8±4.5±2.5	2K	<sup>3</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AB reports $[m_{B_J^+} - m_{B^0}] - m_{\pi^+} = 545.8 \pm 4.1 \pm 2.5$ MeV which we adjust by the $\pi^+$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses two relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>2</sup> AALTONEN 14i reports $m_{B_J(5970)^+} - m_{B^0} - m_{\pi^+} = 541 \pm 5 \pm 12$ MeV which we adjusted by the $\pi^+$ mass.				
<sup>3</sup> AAIJ 15AB reports $[m_{B_J^+} - m_{B^0}] - m_{\pi^+} = 547 \pm 5 \pm 3$ MeV which we adjust by the $\pi^+$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses three relativistic Breit-Wigner functions in the fit for mass difference.				

$m_{B_J(5970)^+} - m_{B^{*0}}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
686.0±4.0±2.5	2k	<sup>4</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>4</sup> AAIJ 15AB reports $[m_{B_J^+} - m_{B^0}] - (m_{B^{*+}} - m_{B^+}) - m_{\pi^+} = 547 \pm 4 \pm 3$ MeV which we adjust by the $\pi^+$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = -(-1)^J$ , $(m_{B^{*0}} - m_{B^0}) = (m_{B^{*+}} - m_{B^+}) = 45.01 \pm 0.30 \pm 0.23$ MeV, and uses three relativistic Breit-Wigner functions in the fit for mass difference.				

$B_J(5970)^+$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>62±20 OUR AVERAGE</b>				
63±15±17	2K	<sup>5</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
60 $^{+30}_{-20}$ ±40	1.4k	AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
61±14±17	2K	<sup>6</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
61±15±17	2K	<sup>7</sup> AAIJ	15AB LHCb	$pp$ at 7, 8 TeV
<sup>5</sup> Assuming $P = (-1)^J$ and using two relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>6</sup> Assuming $P = (-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				
<sup>7</sup> Assuming $P = -(-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				

$B_J(5970)^+$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $B^0\pi^+$	possibly seen	
$\Gamma_2$ $B^{*0}\pi^+$	seen	

Meson Particle Listings

$B_J(5970)^+, B_J(5970)^0$

$B_J(5970)^+$ BRANCHING RATIOS					$B_J(5970)^0$ WIDTH				
$\Gamma(B^0\pi^+)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
possibly seen	2K	<sup>8</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV	<b>81±12 OUR AVERAGE</b>				
possibly seen	1.4k	AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV	82± 8± 9	10K	<sup>5</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV
<sup>8</sup> A $B\pi$ decay is forbidden from a $P = -(-1)^J$ parent, whereas $B^*\pi$ is allowed.					70 <sup>+30</sup> <sub>-20</sub> ±30	2.6k	AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV
					• • • We do not use the following data for averages, fits, limits, etc. • • •				
					56± 7± 9	10K	<sup>6</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV
					82±10± 9	10K	<sup>7</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV
					<sup>5</sup> Assuming $P = (-1)^J$ and using two relativistic Breit-Wigner functions in the fit for mass difference.				
					<sup>6</sup> Assuming $P = (-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				
					<sup>7</sup> Assuming $P = -(-1)^J$ and using three relativistic Breit-Wigner functions in the fit for mass difference.				
$B_J(5970)^+$ REFERENCES					$B_J(5970)^0$ DECAY MODES				
AAIJ	15AB JHEP 1504 024	R. Aaij <i>et al.</i>	(LHCb Collab.)						
AALTONEN	14i PR D90 012013	T. Aaltonen <i>et al.</i>	(CDF Collab.)						
<div><math>B_J(5970)^0</math></div>									
$I(J^P) = \frac{1}{2}(??)$ Status: **									
$I, J, P$ need confirmation.									
Quantum numbers shown are quark-model predictions.									
$B_J(5970)^0$ MASS									
OUR FIT uses $m_{B^+}$ and $m_{B_J(5970)^0} - m_{B^+}$ to determine $m_{B_J(5970)^0}$ .									
VALUE (MeV)		DOCUMENT ID							
<b>5971±5 OUR FIT</b>									
$m_{B_J(5970)^0} - m_{B^+}$									
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT					
<b>691 ±5 OUR FIT</b>									
<b>691 ±5 OUR AVERAGE</b>									
689.9±2.9± 5.1	10K	<sup>1</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV					
698 ±5 ±12	2.6k	<sup>2</sup> AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
714.3±6.4± 5.1	10K	<sup>3</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV					
<sup>1</sup> AAIJ 15AB reports $[m_{B_J^0} - m_{B^+}] - m_{\pi^-} = 550.4 \pm 2.9 \pm 5.1$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses two relativistic Breit-Wigner functions in the fit for mass difference.									
<sup>2</sup> AALTONEN 14i reports $m_{B_J(5970)^0} - m_{B^+} - m_{\pi^-} = 558 \pm 5 \pm 12$ MeV which we adjusted by the $\pi^-$ mass.									
<sup>3</sup> AAIJ 15AB reports $[m_{B_J^0} - m_{B^+}] - m_{\pi^-} = 575 \pm 6 \pm 5$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = (-1)^J$ and uses three relativistic Breit-Wigner functions in the fit for mass difference.									
$m_{B_J(5970)^0} - m_{B^{*+}}$									
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
691.6±3.7±5.1	10k	<sup>4</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV					
<sup>4</sup> AAIJ 15AB reports $[m_{B_J^0} - m_{B^+}] - (m_{B^{*+}} - m_{B^+}) - m_{\pi^-} = 552 \pm 4 \pm 5$ MeV which we adjust by the $\pi^-$ mass. The masses inside the square brackets were measured for each candidate event. The result assumes $P = -(-1)^J$ , $(m_{B^{*+}} - m_{B^+}) = 45.01 \pm 0.30 \pm 0.23$ MeV, and uses three relativistic Breit-Wigner functions in the fit for mass difference.									
$B_J(5970)^0$ BRANCHING RATIOS					$\Gamma_1/\Gamma$				
$\Gamma(B^+\pi^-)/\Gamma_{\text{total}}$									
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT					
possibly seen	10K	<sup>8</sup> AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV					
possibly seen	2.6k	AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV					
<sup>8</sup> A $B\pi$ decay is forbidden from a $P = -(-1)^J$ parent, whereas $B^*\pi$ is allowed.									
$\Gamma(B^{*+}\pi^-)/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT					
seen	10K	AAIJ	15AB LHCb	$p\bar{p}$ at 7, 8 TeV					
seen	2.6k	AALTONEN	14i CDF	$p\bar{p}$ at 1.96 TeV					
$B_J(5970)^0$ REFERENCES									
AAIJ	15AB JHEP 1504 024	R. Aaij <i>et al.</i>	(LHCb Collab.)						
AALTONEN	14i PR D90 012013	T. Aaltonen <i>et al.</i>	(CDF Collab.)						

# BOTTOM, STRANGE MESONS ( $B = \pm 1, S = \mp 1$ )

$$B_s^0 = s\bar{b}, \bar{B}_s^0 = \bar{s}b, \quad \text{similarly for } B_s^{*'}s$$

 $B_s^0$ 

$$I(J^P) = 0(0^-)$$

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

## $B_s^0$ MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>5366.82 ± 0.22 OUR FIT</b>				
<b>5366.7 ± 0.4 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
5366.90 ± 0.28 ± 0.23		<sup>1</sup> AAIJ	12E LHCb	$pp$ at 7 TeV
5364.4 ± 1.3 ± 0.7		<sup>2</sup> LOUVOT	09 BELL	$e^+e^- \rightarrow \Upsilon(5S)$
5366.01 ± 0.73 ± 0.33		<sup>2</sup> ACOSTA	06 CDF	$p\bar{p}$ at 1.96 TeV
5369.9 ± 2.3 ± 1.3	32	<sup>3</sup> ABE	96B CDF	$p\bar{p}$ at 1.8 TeV
5374 ± 16 ± 2	3	<sup>3</sup> ABREU	94D DLPH	$e^+e^- \rightarrow Z$
5359 ± 19 ± 7	1	<sup>3</sup> AKERS	94J OPAL	$e^+e^- \rightarrow Z$
5368.6 ± 5.6 ± 1.5	2	BUSKULIC	93G ALEP	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5370 ± 1 ± 3		DRUTSKOY	07A BELL	Repl. by LOUVOT 09
5370 ± 40	6	<sup>4</sup> AKERS	94J OPAL	$e^+e^- \rightarrow Z$
5383.3 ± 4.5 ± 5.0	14	ABE	93F CDF	Repl. by ABE 96B

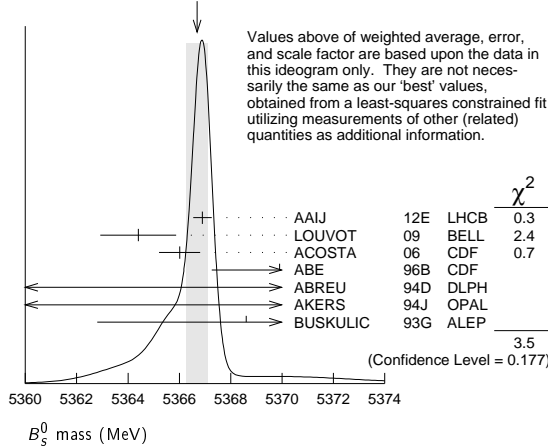
<sup>1</sup> Uses  $B_s^0 \rightarrow J/\psi\phi$  fully reconstructed decays.

<sup>2</sup> Uses exclusively reconstructed final states containing a  $J/\psi \rightarrow \mu^+\mu^-$  decays.

<sup>3</sup> From the decay  $B_s \rightarrow J/\psi(1S)\phi$ .

<sup>4</sup> From the decay  $B_s \rightarrow D_s^-\pi^+$ .

WEIGHTED AVERAGE  
5366.7±0.4 (Error scaled by 1.3)



$$m_{B_s^0} - m_B$$

$m_B$  is the average of our  $B$  masses ( $m_{B^\pm} + m_{B^0}$ )/2.

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>87.35 ± 0.20 OUR FIT</b>				
<b>87.37 ± 0.24 OUR AVERAGE</b>				
87.45 ± 0.44 ± 0.09		<sup>1</sup> AAIJ	15U LHCb	$pp$ at 7, 8 TeV
87.42 ± 0.30 ± 0.09		<sup>2</sup> AAIJ	12E LHCb	$pp$ at 7 TeV
86.64 ± 0.80 ± 0.08		<sup>3</sup> ACOSTA	06 CDF	$p\bar{p}$ at 1.96 TeV
• • • We use the following data for averages but not for fits. • • •				
89.7 ± 2.7 ± 1.2		ABE	96B CDF	$p\bar{p}$ at 1.8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
80 to 130	68	LEE-FRANZINI	90 CSB2	$e^+e^- \rightarrow \Upsilon(5S)$

<sup>1</sup> Uses the mode  $B_s^0 \rightarrow \psi(2S)K^-\pi^+$ .

<sup>2</sup> The reported result is  $m_{B_s^0} - m_{B^+} = 87.52 \pm 0.30 \pm 0.12$  MeV. We convert it to the mass difference with respect to the average of  $(m_{B^\pm} + m_{B^0})/2$ .

<sup>3</sup> The reported result is  $m_{B_s^0} - m_{B^0} = 86.38 \pm 0.90 \pm 0.06$  MeV. We convert it to the mass difference with respect to the average of  $(m_{B^\pm} + m_{B^0})/2$ .

$$m_{B_{sH}^0} - m_{B_{sL}^0}$$

See the  $B_s^0\text{-}\bar{B}_s^0$  MIXING section near the end of these  $B_s^0$  Listings.

## $B_s^0$ MEAN LIFE

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

"OUR EVALUATION" is an average of  $1 / [0.5 (\Gamma_{B_{sL}^0} + \Gamma_{B_{sH}^0})]$ .

VALUE ( $10^{-12}$ s)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1.510 ± 0.005 OUR EVALUATION</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.518 ± 0.041 ± 0.027		<sup>1</sup> AALTONEN	11AP CDF	$p\bar{p}$ at 1.96 TeV
1.398 ± 0.044 ± 0.028		<sup>2</sup> ABAZOV	06V D0	$p\bar{p}$ at 1.96 TeV
1.42 ± 0.14 ± 0.13		<sup>3</sup> ABREU	00Y DLPH	$e^+e^- \rightarrow Z$
1.53 ± 0.16 ± 0.15		<sup>4</sup> ABREU,P	00G DLPH	$e^+e^- \rightarrow Z$
1.36 ± 0.09 ± 0.06		<sup>5</sup> ABE	99D CDF	$p\bar{p}$ at 1.8 TeV
1.72 ± 0.20 ± 0.18		<sup>6</sup> ACKERSTAFF	98F OPAL	$e^+e^- \rightarrow Z$
1.50 ± 0.16 ± 0.15		<sup>5</sup> ACKERSTAFF	98G OPAL	$e^+e^- \rightarrow Z$
1.47 ± 0.14 ± 0.08		<sup>4</sup> BARATE	98C ALEP	$e^+e^- \rightarrow Z$
1.51 ± 0.11		<sup>7</sup> BARATE	98C ALEP	$e^+e^- \rightarrow Z$
1.56 ± 0.29 ± 0.08		<sup>5</sup> ABREU	96F DLPH	Repl. by ABREU 00Y
1.65 ± 0.34 ± 0.31		<sup>4</sup> ABREU	96F DLPH	Repl. by ABREU 00Y
1.76 ± 0.20 ± 0.15		<sup>8</sup> ABREU	96F DLPH	Repl. by ABREU 00Y
1.60 ± 0.26 ± 0.13		<sup>9</sup> ABREU	96F DLPH	Repl. by ABREU,P 00G
1.67 ± 0.14		<sup>10</sup> ABREU	96F DLPH	$e^+e^- \rightarrow Z$
1.61 ± 0.30 ± 0.18	90	<sup>4</sup> BUSKULIC	96E ALEP	Repl. by BARATE 98C
1.54 ± 0.14 ± 0.13		<sup>5</sup> BUSKULIC	96M ALEP	$e^+e^- \rightarrow Z$
1.42 ± 0.27 ± 0.23	76	<sup>5</sup> ABE	95R CDF	Repl. by ABE 99D
1.74 ± 1.08 ± 0.69	8	<sup>11</sup> ABE	95R CDF	Sup. by ABE 96N
1.54 ± 0.25 ± 0.21	79	<sup>5</sup> AKERS	95G OPAL	Repl. by ACKERSTAFF 98G
1.59 ± 0.17 ± 0.15	134	<sup>5</sup> BUSKULIC	95O ALEP	Sup. by BUSKULIC 96M
0.96 ± 0.37	41	<sup>12</sup> ABREU	94E DLPH	Sup. by ABREU 96F
1.92 ± 0.45 ± 0.35	31	<sup>5</sup> BUSKULIC	94C ALEP	Sup. by BUSKULIC 95O
1.13 ± 0.35 ± 0.26	22	<sup>5</sup> ACTON	93H OPAL	Sup. by AKERS 95G

<sup>1</sup> AALTONEN 11AP combines the fully reconstructed  $B_s^0 \rightarrow D_s^-\pi^+$  decays and partially reconstructed  $B_s^0 \rightarrow D_s X$  decays.

<sup>2</sup> Measured using  $D_s\mu^+$  vertices.

<sup>3</sup> Uses  $D_s^-\ell^+$ , and  $\phi\ell^+$  vertices.

<sup>4</sup> Measured using  $D_s$  hadron vertices.

<sup>5</sup> Measured using  $D_s^-\ell^+$  vertices.

<sup>6</sup> ACKERSTAFF 98F use fully reconstructed  $D_s^- \rightarrow \phi\pi^-$  and  $D_s^- \rightarrow K^{*0}K^-$  in the inclusive  $B_s^0$  decay.

<sup>7</sup> Combined results from  $D_s^-\ell^+$  and  $D_s$  hadron.

<sup>8</sup> Measured using  $\phi\ell$  vertices.

<sup>9</sup> Measured using inclusive  $D_s$  vertices.

<sup>10</sup> Combined result for the four ABREU 96F methods.

<sup>11</sup> Exclusive reconstruction of  $B_s \rightarrow \psi\phi$ .

<sup>12</sup> ABREU 94E uses the flight-distance distribution of  $D_s$  vertices,  $\phi$ -lepton vertices, and  $D_s\mu$  vertices.

## $B_s^0$ MEAN LIFE (Flavor specific)

VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
<b>1.511 ± 0.014 OUR EVALUATION</b>			
<b>1.508 ± 0.019 OUR AVERAGE</b>			Error includes scale factor of 1.3. See the ideogram below.
1.479 ± 0.010 ± 0.021	<sup>1</sup> ABAZOV	15A D0	$p\bar{p}$ at 1.96 TeV
1.535 ± 0.015 ± 0.014	<sup>2</sup> AAIJ	14AX LHCb	$pp$ at 7 TeV
1.52 ± 0.15 ± 0.01	<sup>3</sup> AAIJ	14F LHCb	$pp$ at 7, 8 TeV
1.518 ± 0.041 ± 0.027	<sup>4</sup> AALTONEN	11AP CDF	$p\bar{p}$ at 1.96 TeV

## Meson Particle Listings

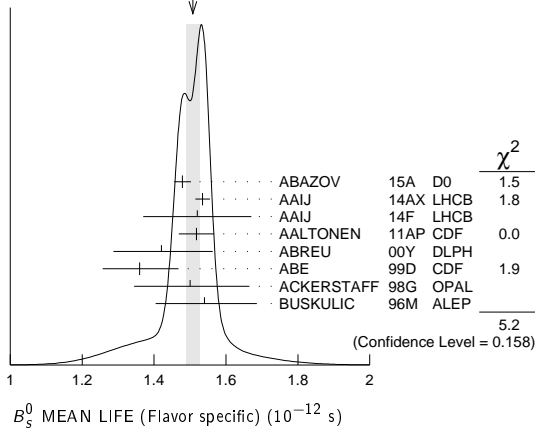
 $B_S^0$ 

1.42 <sup>+0.14</sup> <sub>-0.13</sub> ± 0.03	<sup>5</sup> ABREU	00Y	DLPH	$e^+e^- \rightarrow Z$
1.36 ± 0.09 <sup>+0.06</sup> <sub>-0.05</sub>	<sup>6</sup> ABE	99D	CDF	$p\bar{p}$ at 1.8 TeV
1.50 <sup>+0.16</sup> <sub>-0.15</sub> ± 0.04	<sup>6</sup> ACKERSTAFF	98G	OPAL	$e^+e^- \rightarrow Z$
1.54 <sup>+0.14</sup> <sub>-0.13</sub> ± 0.04	<sup>6</sup> BUSKULIC	96M	ALEP	$e^+e^- \rightarrow Z$
1.60 ± 0.06 ± 0.01	<sup>7</sup> AAIJ	14R	LHCB	$pp$ at 7 TeV
1.398 ± 0.044 <sup>+0.028</sup> <sub>-0.025</sub>	<sup>8</sup> ABAZOV	06V	D0	Repl. by ABAZOV 15A

• • • We do not use the following data for averages, fits, limits, etc. • • •

- Measured using  $B_S^0 \rightarrow D_S^- \mu^+ \nu X$  decays.
- Measured using the  $B_S^0 \rightarrow D_S^- \pi^+$  decays.
- Measured using  $B_S^0 \rightarrow D^+ D_S^-$ .
- AALTONEN 11AP combines the fully reconstructed  $B_S^0 \rightarrow D_S^- \pi^+$  decays and partially reconstructed  $B_S^0 \rightarrow D_S X$  decays.
- Uses  $D_S^- \ell^+$ , and  $\phi \ell^+$  vertices.
- Measured using  $D_S^- \ell^+$  vertices.
- Measured using  $B_S^0 \rightarrow \pi^+ K^-$  decays. May not be flavor specific.
- Measured using  $D_S^- \mu^+$  vertices.

WEIGHTED AVERAGE  
1.508 ± 0.019 (Error scaled by 1.3)

 $B_S^0$  MEAN LIFE ( $B_S \rightarrow J/\psi \phi$ )

VALUE (10 <sup>-12</sup> s)	DOCUMENT ID	TECN	COMMENT
<b>1.479 ± 0.012 OUR EVALUATION</b>			
<b>1.479 ± 0.012 OUR AVERAGE</b>			
1.480 ± 0.011 ± 0.005	<sup>1</sup> AAIJ	14E	LHCB $pp$ at 7 TeV
1.444 <sup>+0.098</sup> <sub>-0.096</sub> ± 0.020	<sup>1</sup> ABAZOV	05B	D0 $p\bar{p}$ at 1.96 TeV
1.34 <sup>+0.23</sup> <sub>-0.19</sub> ± 0.05	<sup>2</sup> ABE	98B	CDF $p\bar{p}$ at 1.8 TeV
1.39 <sup>+0.13</sup> <sub>-0.16</sub> ± 0.02	<sup>2</sup> ABAZOV	05W	D0 $p\bar{p}$ at 1.96 TeV
1.34 <sup>+0.23</sup> <sub>-0.19</sub> ± 0.05	<sup>3</sup> ABE	96N	CDF Repl. by ABE 98B

• • • We do not use the following data for averages, fits, limits, etc. • • •

- Measured using fully reconstructed  $B_S \rightarrow J/\psi \phi$  decays.
- Measured using the time-dependent angular analysis of  $B_S^0 \rightarrow J/\psi \phi$  decays.
- ABE 96N uses 58 ± 12 exclusive  $B_S \rightarrow J/\psi \phi$  events.

 $B_{sH}^0$  MEAN LIFE

$B_{sH}^0$  is the heavy mass state of two  $B_S^0$  CP eigenstates.

“OUR EVALUATION” has been obtained by the Heavy Flavor Averaging Group (HFAG) using the constraint of the flavor-specific lifetime average in a way similar to  $\Delta\Gamma_{B_S^0}/\Gamma_{B_S^0}$ .

VALUE (10 <sup>-12</sup> s)	DOCUMENT ID	TECN	COMMENT
<b>1.610 ± 0.012 OUR EVALUATION</b>			
<b>1.70 ± 0.04 OUR AVERAGE</b>			
1.75 ± 0.12 ± 0.07	<sup>1</sup> AAIJ	13AB	LHCB $pp$ at 7 TeV
1.700 ± 0.040 ± 0.026	<sup>2</sup> AAIJ	12AN	LHCB $pp$ at 7 TeV
1.70 <sup>+0.12</sup> <sub>-0.11</sub> ± 0.03	<sup>2</sup> AALTONEN	11AB	CDF $p\bar{p}$ at 1.96 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.613 <sup>+0.123</sup> <sub>-0.113</sub>	<sup>3</sup> AALTONEN	12D	CDF $p\bar{p}$ at 1.96 TeV
1.58 <sup>+0.39</sup> <sub>-0.42</sub> ± 0.01	<sup>4,5</sup> AALTONEN	08J	CDF Repl. by AALTONEN 12D
2.07 <sup>+0.58</sup> <sub>-0.46</sub> ± 0.03	<sup>5</sup> ABAZOV	05W	D0 Repl. by ABAZOV 08AM
	<sup>5</sup> ACOSTA	05	CDF Repl. by AALTONEN 08J

- Measured using a pure CP-odd final state  $J/\psi K_S^0$  with the assumption that contributions from penguin diagrams are small.
- Measured using a pure CP-odd final state  $J/\psi f_0(980)$ .
- Uses the time-dependent angular analysis of  $B_S^0 \rightarrow J/\psi \phi$  decays assuming CP-violating angle  $\beta_S(B^0 \rightarrow J/\psi \phi) = 0.02$ .
- Obtained from  $\Delta\Gamma_S$  and  $\Gamma_S$  fit with a correlation of 0.6.
- Measured using the time-dependent angular analysis of  $B_S^0 \rightarrow J/\psi \phi$  decays.

 $B_{sL}^0$  MEAN LIFE

$B_{sL}^0$  is the light mass state of two  $B_S^0$  CP eigenstates.

“OUR EVALUATION” has been obtained by the Heavy Flavor Averaging Group (HFAG) using the constraint of the flavor-specific lifetime average in a way similar to  $\Delta\Gamma_{B_S^0}/\Gamma_{B_S^0}$ .

VALUE (10 <sup>-12</sup> s)	DOCUMENT ID	TECN	COMMENT
<b>1.422 ± 0.008 OUR EVALUATION</b>			
<b>1.379 ± 0.026 ± 0.017</b>	<sup>1</sup> AAIJ	14F	LHCB $pp$ at 7, 8 TeV
1.407 ± 0.016 ± 0.007	<sup>2</sup> AAIJ	14R	LHCB $pp$ at 7 TeV
1.440 ± 0.096 ± 0.009	<sup>2</sup> AAIJ	12	LHCB Repl. by AAIJ 14R
1.455 ± 0.046 ± 0.006	<sup>2</sup> AAIJ	12R	LHCB Repl. by AAIJ 14R
1.437 <sup>+0.054</sup> <sub>-0.047</sub>	<sup>3</sup> AALTONEN	12D	CDF $p\bar{p}$ at 1.96 TeV
1.24 <sup>+0.14</sup> <sub>-0.11</sub> ± 0.01	<sup>4,5</sup> AALTONEN	08J	CDF Repl. by AALTONEN 12D
1.05 <sup>+0.16</sup> <sub>-0.13</sub> ± 0.02	<sup>5</sup> ABAZOV	05W	D0 Repl. by ABAZOV 08AM
1.27 ± 0.33 ± 0.08	<sup>5</sup> ACOSTA	05	CDF Repl. by AALTONEN 08J
	<sup>6</sup> BARATE	00K	ALEP $e^+e^- \rightarrow Z$

- Measured using  $B_S^0 \rightarrow D_S^- D_S^+$ . The effective lifetime is translated into a decay width of  $\Gamma_L = 0.725 \pm 0.014 \pm 0.009 \text{ ps}^{-1}$ .
- Measured using  $B_S^0 \rightarrow K^+ K^-$  decays. There may still be CPV in the decay.
- Uses the time-dependent angular analysis of  $B_S^0 \rightarrow J/\psi \phi$  decays and assuming CP-violating angle  $\beta_S(B^0 \rightarrow J/\psi \phi) = 0.02$ .
- Obtained from  $\Delta\Gamma_S$  and  $\Gamma_S$  fit with a correlation of 0.6.
- Measured using the time-dependent angular analysis of  $B_S^0 \rightarrow J/\psi \phi$  decays.
- Uses  $\phi \phi$  correlations from  $B_S^0 \rightarrow D_S^{(*)+} D_S^{(*)-}$ .

 $\Delta\Gamma_{B_S^0}/\Gamma_{B_S^0}$ 

$\Gamma_{B_S^0}$  and  $\Delta\Gamma_{B_S^0}$  are the decay rate average and difference between two  $B_S^0$  CP eigenstates (light – heavy).

“OUR EVALUATION” is an average of all available  $B_S$  flavor-specific lifetime measurements with the  $\Delta\Gamma_{B_S^0}/\Gamma_S$  analyses performed by the Heavy

Flavor Averaging Group (HFAG) as described in our “Review on  $B-\bar{B}$  Mixing” in the  $B^0$  Section of these Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.124 ± 0.011 OUR EVALUATION</b>				
0.090 ± 0.009 ± 0.023		<sup>1</sup> AAIJ	12D	LHCB $pp$ at 7 TeV
		<sup>2</sup> ABAZOV	12D	D0 $p\bar{p}$ at 1.96 TeV
0.147 <sup>+0.036</sup> <sub>-0.030</sub> ± 0.042		<sup>3</sup> ESEN	13	BELL $e^+e^- \rightarrow \gamma(5S)$
0.116 <sup>+0.09</sup> <sub>-0.10</sub> ± 0.010		<sup>4</sup> AALTONEN	12D	CDF $p\bar{p}$ at 1.96 TeV
0.24 <sup>+0.28</sup> <sub>-0.38</sub> ± 0.03		<sup>3</sup> ESEN	10	BELL $e^+e^- \rightarrow \gamma(5S)$
0.65 <sup>+0.25</sup> <sub>-0.33</sub> ± 0.01		<sup>5</sup> AALTONEN	08J	CDF Repl. by AALTONEN 12D
		<sup>5,6</sup> ABAZOV	05W	D0 Repl. by ABAZOV 08AM
		<sup>5</sup> ACOSTA	05	CDF Repl. by AALTONEN 08J
<0.46	95	<sup>7</sup> ABREU	00Y	DLPH $e^+e^- \rightarrow Z$
<0.69	95	<sup>8</sup> ABREU,P	00G	DLPH $e^+e^- \rightarrow Z$
<0.83	95	<sup>9</sup> ABE	99D	CDF $p\bar{p}$ at 1.8 TeV
<0.67	95	<sup>10</sup> ACCIARRI	98S	L3 $e^+e^- \rightarrow Z$

See key on page 601

## Meson Particle Listings

 $B_s^0$ 

- <sup>1</sup> Measured using the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>2</sup> Measured using fully reconstructed  $B_s \rightarrow J/\psi\phi$  decays.  
<sup>3</sup> Assumes  $CP$  violation is negligible.  
<sup>4</sup> Uses the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays and assuming  $CP$ -violating angle  $\beta_s(B^0 \rightarrow J/\psi\phi) = 0.02$ .  
<sup>5</sup> Measured using the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>6</sup> Uses  $|A_0|^2 - |A_{||}|^2 = 0.355 \pm 0.066$  from ACOSTA 05.  
<sup>7</sup> Uses  $D_s^- \ell^+$  and  $\phi \ell^+$  vertices.  
<sup>8</sup> Measured using  $D_s$  hadron vertices.  
<sup>9</sup> ABE 99d assumes  $\tau_{B_s^0} = 1.55 \pm 0.05$  ps.  
<sup>10</sup> ACCIARRI 98s assumes  $\tau_{B_s^0} = 1.49 \pm 0.06$  ps and PDG 98 values of  $b$  production fraction.

 $\Delta\Gamma_{B_s^0}$ 

"OUR EVALUATION" has been obtained by the Heavy Flavor Averaging Group (HFAG) using the constraint of the flavor-specific lifetime average in a way similar to  $\Delta\Gamma_{B_s^0}/\Gamma_{B_s^0}$ .

VALUE ( $10^{12} \text{ s}^{-1}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.082 <math>\pm 0.007</math> OUR EVALUATION</b>			
<b>0.077 <math>\pm 0.008</math> OUR AVERAGE</b>			
0.0805 $\pm 0.0091 \pm 0.0032$	<sup>1</sup> AAIJ	15i LHCb	$pp$ at 7, 8 TeV
0.053 $\pm 0.021 \pm 0.010$	<sup>2</sup> AAD	14u ATLS	$pp$ at 7 TeV
0.068 $\pm 0.026 \pm 0.007$	<sup>3</sup> AALTONEN	12AJ CDF	$p\bar{p}$ at 1.96 TeV
0.163 $^{+0.065}_{-0.064}$	<sup>4,5</sup> ABAZOV	12D D0	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.106 $\pm 0.011 \pm 0.007$	<sup>6</sup> AAIJ	13AR LHCb	Repl. by AAIJ 15i
0.053 $\pm 0.021 \pm 0.010$	<sup>3</sup> AAD	12CV ATLS	Repl. by AAD 14u
0.123 $\pm 0.029 \pm 0.011$	<sup>3</sup> AAIJ	12D LHCb	Repl. by AAIJ 13AR
0.075 $\pm 0.035 \pm 0.006$	<sup>7</sup> AALTONEN	12D CDF	Repl. by AALTONEN 12AJ
0.085 $^{+0.072}_{-0.078} \pm 0.001$	<sup>8</sup> ABAZOV	09E D0	Repl. by ABAZOV 08AM
0.076 $^{+0.059}_{-0.063} \pm 0.006$	<sup>9</sup> AALTONEN	08J CDF	Repl. by AALTONEN 12D
0.19 $\pm 0.07$ $^{+0.02}_{-0.01}$	<sup>5,10</sup> ABAZOV	08AMD0	Repl. by ABAZOV 12D
0.12 $^{+0.08}_{-0.10} \pm 0.02$	<sup>9,11</sup> ABAZOV	07 D0	Repl. by ABAZOV 07N
0.13 $\pm 0.09$	<sup>12</sup> ABAZOV	07N D0	Repl. by ABAZOV 09E
0.47 $^{+0.19}_{-0.24} \pm 0.01$	<sup>9</sup> ACOSTA	05 CDF	Repl. by AALTONEN 08J

- <sup>1</sup> Measured using time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi K^+ K^-$  decays.  
<sup>2</sup> Measured using the flavor tagged time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>3</sup> Measured using the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>4</sup> The error includes both statistical and systematic uncertainties.  
<sup>5</sup> Measured using fully reconstructed  $B_s \rightarrow J/\psi\phi$  decays.  
<sup>6</sup> AAIJ 13AR result comes from a combined fit to  $B_s^0 \rightarrow J/\psi K^+ K^-$  and  $B_s^0 \rightarrow J/\psi\pi^+\pi^-$  data sets. Also reports  $\Delta\Gamma_s = 0.100 \pm 0.016 \pm 0.003 \text{ ps}^{-1}$  from a fit to  $B_s^0 \rightarrow J/\psi K^+ K^-$  decays.  
<sup>7</sup> Uses the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays and assuming  $CP$ -violating angle  $\beta_s(B^0 \rightarrow J/\psi\phi) = 0.02$ .  
<sup>8</sup> Measured the angular and lifetime parameters for the time-dependent angular untagged decays  $B_d^0 \rightarrow J/\psi K^{*0}$  and  $B_s^0 \rightarrow J/\psi\phi$ .  
<sup>9</sup> Measured using the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays and assuming  $CP$ -violating phase  $\phi_s = 0$ .  
<sup>10</sup> Obtains 90% CL interval  $-0.06 < \Delta\Gamma_s < 0.30$ .  
<sup>11</sup> ABAZOV 07 reports  $0.17 \pm 0.09 \pm 0.02$  with  $CP$ -violating phase  $\phi_s$  as a free parameter.  
<sup>12</sup> Combines  $D^0$  measurements of time-dependent angular distributions in  $B_s^0 \rightarrow J/\psi\phi$  and charge asymmetry in semileptonic decays. There is a 4-fold ambiguity in the solution.

 $\Delta\Gamma_s^{CP} / \Gamma_s$ 

$\Gamma_s$  and  $\Delta\Gamma_s^{CP}$  are the decay rate average and difference between even,  $\Gamma_s^{CP\text{-even}}$ , and odd,  $\Gamma_s^{CP\text{-odd}}$ ,  $CP$  eigenstates.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.072 $\pm 0.021 \pm 0.022$		<sup>1</sup> ABAZOV	09i D0	$p\bar{p}$ at 1.96 TeV
>0.012	95	<sup>1</sup> AALTONEN	08F CDF	$p\bar{p}$ at 1.96 TeV
0.079 $^{+0.038}_{-0.035} \pm 0.031$		<sup>1</sup> ABAZOV	07Y D0	Repl. by ABAZOV 09i
0.25 $^{+0.21}_{-0.14}$		<sup>2</sup> BARATE	00K ALEP	$e^+ e^- \rightarrow Z$

- <sup>1</sup> Assumes  $2 \text{ B}(B_s^0 \rightarrow D_s^{(*)} D_s^{(*)}) \simeq \Delta\Gamma_s^{CP} / \Gamma_s$ .  
<sup>2</sup> Uses  $\phi\phi$  correlations from  $B_s^0 \rightarrow D_s^{(*)+} D_s^{(*)-}$ .

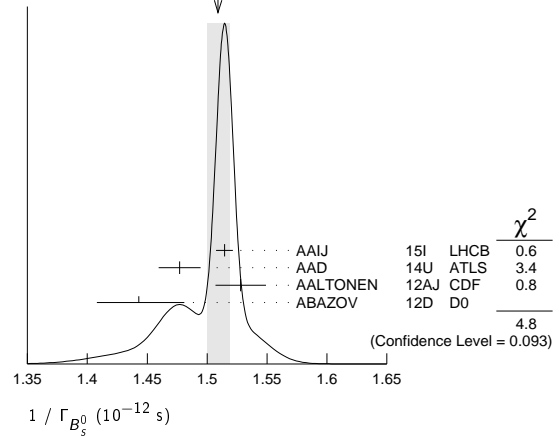
 $1 / \Gamma_{B_s^0}$ 

"OUR EVALUATION" has been obtained by the Heavy Flavor Averaging Group (HFAG) using the constraint of the flavor-specific lifetime average in a way similar to  $\Delta\Gamma_{B_s^0}/\Gamma_{B_s^0}$ .

VALUE ( $10^{-12} \text{ s}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.510 <math>\pm 0.005</math> OUR EVALUATION</b>			
<b>1.509 <math>\pm 0.010</math> OUR AVERAGE</b>			Error includes scale factor of 1.5. See the ideogram below.
1.5145 $\pm 0.0062 \pm 0.0034$	<sup>1</sup> AAIJ	15i LHCb	$pp$ at 7, 8 TeV
1.477 $\pm 0.015 \pm 0.009$	<sup>2</sup> AAD	14u ATLS	$pp$ at 7 TeV
1.528 $\pm 0.019 \pm 0.009$	<sup>3</sup> AALTONEN	12AJ CDF	$p\bar{p}$ at 1.96 TeV
1.443 $^{+0.038}_{-0.035}$	<sup>3,4</sup> ABAZOV	12D D0	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.513 $\pm 0.009 \pm 0.014$	<sup>5</sup> AAIJ	13AR LHCb	Repl. by AAIJ 15i
1.477 $\pm 0.015 \pm 0.009$	<sup>6</sup> AAD	12CV ATLS	Repl. by AAD 14u
1.522 $\pm 0.021 \pm 0.019$	<sup>7</sup> AAIJ	12D LHCb	Repl. by AAIJ 13AR
1.529 $\pm 0.025 \pm 0.012$	<sup>3</sup> AALTONEN	12D CDF	Repl. by AALTONEN 12AJ
1.487 $\pm 0.060 \pm 0.028$	<sup>3</sup> ABAZOV	09E D0	Repl. by ABAZOV 08AM
1.52 $\pm 0.04 \pm 0.02$	<sup>3</sup> AALTONEN	08J CDF	Repl. by AALTONEN 12D
1.52 $\pm 0.05 \pm 0.01$	<sup>3</sup> ABAZOV	08AMD0	Repl. by ABAZOV 12D
1.40 $^{+0.15}_{-0.13} \pm 0.02$	<sup>3</sup> ACOSTA	05 CDF	$p\bar{p}$ at 1.96 TeV

- <sup>1</sup> AAIJ 15i reports  $\Gamma_{B_s^0} = 0.6603 \pm 0.0027 \pm 0.0015 \text{ ps}^{-1}$  obtained from time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi K^+ K^-$  decays.  
<sup>2</sup> AAD 14u reports  $\Gamma_{B_s^0} = 0.677 \pm 0.007 \pm 0.004 \text{ ps}^{-1}$  measured using a tagged, time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>3</sup> Measured using the time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>4</sup> The error includes both statistical and systematic uncertainties.  
<sup>5</sup> AAIJ 13AR reports  $\Gamma_s = 0.661 \pm 0.004 \pm 0.006 \text{ ps}^{-1}$  obtained from combined fit to  $B_s^0 \rightarrow J/\psi K^+ K^-$  and  $B_s^0 \rightarrow J/\psi\pi^+\pi^-$  data sets. Also reports a separate measurement of  $\Gamma_s = 0.663 \pm 0.005 \pm 0.006 \text{ ps}^{-1}$  from  $B_s^0 \rightarrow J/\psi K^+ K^-$  decays.  
<sup>6</sup> AAD 12CV reports  $\Gamma_{B_s^0} = 0.677 \pm 0.007 \pm 0.004 \text{ ps}^{-1}$  measured using a time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi\phi$  decays.  
<sup>7</sup> AAIJ 12D reports average decay width of  $B_s^0$ ,  $\Gamma_{B_s^0} = 0.657 \pm 0.009 \pm 0.008 \text{ ps}^{-1}$  that we converted to  $1/\Gamma_{B_s^0}$ .

WEIGHTED AVERAGE  
1.509 $\pm 0.010$  (Error scaled by 1.5)

 $B_s^0$  DECAY MODES

These branching fractions all scale with  $\text{B}(\bar{b} \rightarrow B_s^0)$ .

The branching fraction  $\text{B}(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything})$  is not a pure measurement since the measured product branching fraction  $\text{B}(\bar{b} \rightarrow B_s^0) \times \text{B}(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything})$  was used to determine  $\text{B}(\bar{b} \rightarrow B_s^0)$ , as described in the note on " $B^0$ - $\bar{B}^0$  Mixing"

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm \text{ anything}$ , the values usually are multiplicities, not branching fractions. They can be greater than one.

## Meson Particle Listings

 $B_s^0$ 

	Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$	$D_s^-$ anything	(93 $\pm$ 25 ) %	
$\Gamma_2$	$\ell \nu_\ell X$	( 9.6 $\pm$ 0.8 ) %	
$\Gamma_3$	$e^+ \nu X^-$	( 9.1 $\pm$ 0.8 ) %	
$\Gamma_4$	$\mu^+ \nu X^-$	(10.2 $\pm$ 1.0 ) %	
$\Gamma_5$	$D_s^- \ell^+ \nu_\ell$ anything	[a] ( 8.1 $\pm$ 1.3 ) %	
$\Gamma_6$	$D_s^{*-} \ell^+ \nu_\ell$ anything	( 5.4 $\pm$ 1.1 ) %	
$\Gamma_7$	$D_{s1}(2536)^- \mu^+ \nu_\mu$	( 2.6 $\pm$ 0.7 ) $\times 10^{-3}$	
$\Gamma_8$	$D_{s1}^- \rightarrow D^{*-} K_S^0$ $D_{s1}^-(2536)^- X \mu^+ \nu$	( 4.4 $\pm$ 1.3 ) $\times 10^{-3}$	
$\Gamma_9$	$D_{s2}(2573)^- X \mu^+ \nu$ $D_{s2}^- \rightarrow \bar{D}^0 K^+$	( 2.7 $\pm$ 1.0 ) $\times 10^{-3}$	
$\Gamma_{10}$	$D_s^- \pi^+$	( 3.00 $\pm$ 0.23 ) $\times 10^{-3}$	
$\Gamma_{11}$	$D_s^- \rho^+$	( 6.9 $\pm$ 1.4 ) $\times 10^{-3}$	
$\Gamma_{12}$	$D_s^- \pi^+ \pi^+ \pi^-$	( 6.1 $\pm$ 1.0 ) $\times 10^{-3}$	
$\Gamma_{13}$	$D_{s1}(2536)^- \pi^+$ $D_{s1}^- \rightarrow D_s^- \pi^+ \pi^-$	( 2.5 $\pm$ 0.8 ) $\times 10^{-5}$	
$\Gamma_{14}$	$D_s^\mp K^\pm$	( 2.27 $\pm$ 0.19 ) $\times 10^{-4}$	
$\Gamma_{15}$	$D_s^- K^+ \pi^+ \pi^-$	( 3.2 $\pm$ 0.6 ) $\times 10^{-4}$	
$\Gamma_{16}$	$D_s^+ D_s^-$	( 4.4 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{17}$	$D_s^- D^+$	( 2.8 $\pm$ 0.5 ) $\times 10^{-4}$	
$\Gamma_{18}$	$D^+ D^-$	( 2.2 $\pm$ 0.6 ) $\times 10^{-4}$	
$\Gamma_{19}$	$D^0 \bar{D}^0$	( 1.9 $\pm$ 0.5 ) $\times 10^{-4}$	
$\Gamma_{20}$	$D_s^{*-} \pi^+$	( 2.0 $\pm$ 0.5 ) $\times 10^{-3}$	
$\Gamma_{21}$	$D_s^{*\mp} K^\pm$	( 1.33 $\pm$ 0.35 ) $\times 10^{-4}$	
$\Gamma_{22}$	$D_s^{*-} \rho^+$	( 9.6 $\pm$ 2.1 ) $\times 10^{-3}$	
$\Gamma_{23}$	$D_s^{*+} D_s^- + D_s^{*-} D_s^+$	( 1.29 $\pm$ 0.22 ) %	S=1.1
$\Gamma_{24}$	$D_s^{*+} D_s^-$	( 1.86 $\pm$ 0.30 ) %	
$\Gamma_{25}$	$D_s^{(*)+} D_s^{(*)-}$	( 4.5 $\pm$ 1.4 ) %	
$\Gamma_{26}$	$\bar{D}^0 K^- \pi^+$	( 1.03 $\pm$ 0.13 ) $\times 10^{-3}$	
$\Gamma_{27}$	$\bar{D}^0 K^*(892)^0$	( 4.4 $\pm$ 0.6 ) $\times 10^{-4}$	
$\Gamma_{28}$	$\bar{D}^0 K^*(1410)$	( 3.9 $\pm$ 3.5 ) $\times 10^{-4}$	
$\Gamma_{29}$	$\bar{D}^0 K_0^*(1430)$	( 3.0 $\pm$ 0.7 ) $\times 10^{-4}$	
$\Gamma_{30}$	$\bar{D}^0 K_2^*(1430)$	( 1.1 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{31}$	$\bar{D}^0 K^*(1680)$	< 7.8 $\times 10^{-5}$	CL=90%
$\Gamma_{32}$	$\bar{D}^0 K_0^*(1950)$	< 1.1 $\times 10^{-4}$	CL=90%
$\Gamma_{33}$	$\bar{D}^0 K_2^*(1780)$	< 2.6 $\times 10^{-5}$	CL=90%
$\Gamma_{34}$	$\bar{D}^0 K_4^*(2045)$	< 3.1 $\times 10^{-5}$	CL=90%
$\Gamma_{35}$	$\bar{D}^0 K^- \pi^+$ (non-resonant)	( 2.1 $\pm$ 0.8 ) $\times 10^{-4}$	
$\Gamma_{36}$	$D_{s2}^*(2573)^- \pi^+$ $D_{s2}^* \rightarrow \bar{D}^0 K^-$	( 2.6 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{37}$	$D_{s1}^*(2700)^- \pi^+$ $D_{s1}^* \rightarrow \bar{D}^0 K^-$	( 1.6 $\pm$ 0.8 ) $\times 10^{-5}$	
$\Gamma_{38}$	$D_{s1}^*(2860)^- \pi^+$ $D_{s1}^* \rightarrow \bar{D}^0 K^-$	( 5 $\pm$ 4 ) $\times 10^{-5}$	
$\Gamma_{39}$	$D_{s3}^*(2860)^- \pi^+$ $D_{s3}^* \rightarrow \bar{D}^0 K^-$	( 2.2 $\pm$ 0.6 ) $\times 10^{-5}$	
$\Gamma_{40}$	$\bar{D}^0 K^+ K^-$	( 4.4 $\pm$ 2.0 ) $\times 10^{-5}$	
$\Gamma_{41}$	$\bar{D}^0 f_0(980)$	< 3.1 $\times 10^{-6}$	CL=90%
$\Gamma_{42}$	$\bar{D}^0 \phi$	( 3.0 $\pm$ 0.8 ) $\times 10^{-5}$	
$\Gamma_{43}$	$D^{*\mp} \pi^\pm$	< 6.1 $\times 10^{-6}$	CL=90%
$\Gamma_{44}$	$J/\psi(1S) \phi$	( 1.07 $\pm$ 0.08 ) $\times 10^{-3}$	
$\Gamma_{45}$	$J/\psi(1S) \pi^0$	< 1.2 $\times 10^{-3}$	CL=90%
$\Gamma_{46}$	$J/\psi(1S) \eta$	( 3.9 $\pm$ 0.7 ) $\times 10^{-4}$	S=1.4
$\Gamma_{47}$	$J/\psi(1S) K_S^0$	( 1.89 $\pm$ 0.12 ) $\times 10^{-5}$	
$\Gamma_{48}$	$J/\psi(1S) \bar{K}^*(892)^0$	( 4.1 $\pm$ 0.4 ) $\times 10^{-5}$	
$\Gamma_{49}$	$J/\psi(1S) \eta'$	( 3.3 $\pm$ 0.4 ) $\times 10^{-4}$	
$\Gamma_{50}$	$J/\psi(1S) \pi^+ \pi^-$	( 2.13 $\pm$ 0.18 ) $\times 10^{-4}$	
$\Gamma_{51}$	$J/\psi(1S) f_0(500), f_0 \rightarrow \pi^+ \pi^-$	< 1.7 $\times 10^{-6}$	CL=90%
$\Gamma_{52}$	$J/\psi(1S) \rho, \rho \rightarrow \pi^+ \pi^-$	< 1.2 $\times 10^{-6}$	CL=90%
$\Gamma_{53}$	$J/\psi(1S) f_0(980), f_0 \rightarrow \pi^+ \pi^-$	( 1.34 $\pm$ 0.15 ) $\times 10^{-4}$	
$\Gamma_{54}$	$J/\psi(1S) f_0(980)_0, f_0 \rightarrow \pi^+ \pi^-$	( 5.1 $\pm$ 0.9 ) $\times 10^{-5}$	

$\Gamma_{55}$	$J/\psi(1S) f_2(1270), f_2 \rightarrow \pi^+ \pi^-$		
$\Gamma_{56}$	$J/\psi(1S) f_2(1270)_0, f_2 \rightarrow \pi^+ \pi^-$	( 2.6 $\pm$ 0.7 ) $\times 10^{-7}$	
$\Gamma_{57}$	$J/\psi(1S) f_2(1270)_\parallel, f_2 \rightarrow \pi^+ \pi^-$	( 3.8 $\pm$ 1.3 ) $\times 10^{-7}$	
$\Gamma_{58}$	$J/\psi(1S) f_2(1270)_\perp, f_2 \rightarrow \pi^+ \pi^-$	( 4.6 $\pm$ 2.7 ) $\times 10^{-7}$	
$\Gamma_{59}$	$J/\psi(1S) f_0(1370), f_0 \rightarrow \pi^+ \pi^-$		
$\Gamma_{60}$	$J/\psi(1S) f_0(1500), f_0 \rightarrow \pi^+ \pi^-$	( 7.3 $\pm$ 1.6 ) $\times 10^{-6}$	
$\Gamma_{61}$	$J/\psi(1S) f'_2(1525)_0, f'_2 \rightarrow \pi^+ \pi^-$	( 3.7 $\pm$ 1.0 ) $\times 10^{-7}$	
$\Gamma_{62}$	$J/\psi(1S) f'_2(1525)_\parallel, f'_2 \rightarrow \pi^+ \pi^-$	( 4.3 $\pm$ 9.0 ) $\times 10^{-8}$	
$\Gamma_{63}$	$J/\psi(1S) f'_2(1525)_\perp, f'_2 \rightarrow \pi^+ \pi^-$	( 1.9 $\pm$ 1.4 ) $\times 10^{-7}$	
$\Gamma_{64}$	$J/\psi(1S) f_0(1790), f_0 \rightarrow \pi^+ \pi^-$	( 1.7 $\pm$ 4.0 ) $\times 10^{-6}$	
$\Gamma_{65}$	$J/\psi(1S) \pi^+ \pi^-$ (nonresonant)		
$\Gamma_{66}$	$J/\psi(1S) \bar{K}^0 \pi^+ \pi^-$	< 4.4 $\times 10^{-5}$	CL=90%
$\Gamma_{67}$	$J/\psi(1S) K^+ K^-$	( 7.9 $\pm$ 0.7 ) $\times 10^{-4}$	
$\Gamma_{68}$	$J/\psi(1S) K^0 K^- \pi^+ + \text{c.c.}$	( 9.3 $\pm$ 1.3 ) $\times 10^{-4}$	
$\Gamma_{69}$	$J/\psi(1S) \bar{K}^0 K^+ K^-$	< 1.2 $\times 10^{-5}$	CL=90%
$\Gamma_{70}$	$J/\psi(1S) f'_2(1525)$	( 2.6 $\pm$ 0.6 ) $\times 10^{-4}$	
$\Gamma_{71}$	$J/\psi(1S) \rho \bar{\rho}$	< 4.8 $\times 10^{-6}$	CL=90%
$\Gamma_{72}$	$J/\psi(1S) \gamma$	< 7.3 $\times 10^{-6}$	CL=90%
$\Gamma_{73}$	$J/\psi(1S) \pi^+ \pi^- \pi^+ \pi^-$	( 7.9 $\pm$ 0.9 ) $\times 10^{-5}$	
$\Gamma_{74}$	$J/\psi(1S) f_1(1285)$	( 7.1 $\pm$ 1.4 ) $\times 10^{-5}$	
$\Gamma_{75}$	$\psi(2S) \eta$	( 3.3 $\pm$ 0.9 ) $\times 10^{-4}$	
$\Gamma_{76}$	$\psi(2S) \eta'$	( 1.29 $\pm$ 0.35 ) $\times 10^{-4}$	
$\Gamma_{77}$	$\psi(2S) \pi^+ \pi^-$	( 7.2 $\pm$ 1.2 ) $\times 10^{-5}$	
$\Gamma_{78}$	$\psi(2S) \phi$	( 5.4 $\pm$ 0.5 ) $\times 10^{-4}$	
$\Gamma_{79}$	$\psi(2S) K^- \pi^+$	( 3.12 $\pm$ 0.30 ) $\times 10^{-5}$	
$\Gamma_{80}$	$\psi(2S) \bar{K}^*(892)^0$	( 3.3 $\pm$ 0.5 ) $\times 10^{-5}$	
$\Gamma_{81}$	$\chi_{c1} \phi$	( 2.03 $\pm$ 0.29 ) $\times 10^{-4}$	
$\Gamma_{82}$	$\pi^+ \pi^-$	( 7.7 $\pm$ 2.0 ) $\times 10^{-7}$	S=1.4
$\Gamma_{83}$	$\pi^0 \pi^0$	< 2.1 $\times 10^{-4}$	CL=90%
$\Gamma_{84}$	$\eta \pi^0$	< 1.0 $\times 10^{-3}$	CL=90%
$\Gamma_{85}$	$\eta \eta$	< 1.5 $\times 10^{-3}$	CL=90%
$\Gamma_{86}$	$\rho^0 \rho^0$	< 3.20 $\times 10^{-4}$	CL=90%
$\Gamma_{87}$	$\eta' \eta'$	( 3.3 $\pm$ 0.7 ) $\times 10^{-5}$	
$\Gamma_{88}$	$\phi \rho^0$	< 6.17 $\times 10^{-4}$	CL=90%
$\Gamma_{89}$	$\phi \phi$	( 1.87 $\pm$ 0.15 ) $\times 10^{-5}$	
$\Gamma_{90}$	$\pi^+ K^-$	( 5.6 $\pm$ 0.6 ) $\times 10^{-6}$	
$\Gamma_{91}$	$K^+ K^-$	( 2.52 $\pm$ 0.17 ) $\times 10^{-5}$	
$\Gamma_{92}$	$K^0 \bar{K}^0$	< 6.6 $\times 10^{-5}$	CL=90%
$\Gamma_{93}$	$K^0 \pi^+ \pi^-$	( 1.5 $\pm$ 0.4 ) $\times 10^{-5}$	
$\Gamma_{94}$	$K^0 K^\pm \pi^\mp$	( 7.7 $\pm$ 1.0 ) $\times 10^{-5}$	
$\Gamma_{95}$	$K^*(892)^- \pi^+$	( 3.3 $\pm$ 1.2 ) $\times 10^{-6}$	
$\Gamma_{96}$	$K^*(892)^\pm K^\mp$	( 1.25 $\pm$ 0.26 ) $\times 10^{-5}$	
$\Gamma_{97}$	$K_S^0 \bar{K}^*(892)^0 + \text{c.c.}$	( 1.6 $\pm$ 0.4 ) $\times 10^{-5}$	
$\Gamma_{98}$	$K^0 K^+ K^-$	< 3.5 $\times 10^{-6}$	CL=90%
$\Gamma_{99}$	$\bar{K}^*(892)^0 \rho^0$	< 7.67 $\times 10^{-4}$	CL=90%
$\Gamma_{100}$	$\bar{K}^*(892)^0 K^*(892)^0$	( 1.11 $\pm$ 0.27 ) $\times 10^{-5}$	
$\Gamma_{101}$	$\phi K^*(892)^0$	( 1.14 $\pm$ 0.30 ) $\times 10^{-6}$	
$\Gamma_{102}$	$\rho \bar{\rho}$	( 2.8 $\pm$ 2.2 ) $\times 10^{-8}$	
$\Gamma_{103}$	$\Lambda_c^- \Lambda \pi^+$	( 3.6 $\pm$ 1.6 ) $\times 10^{-4}$	
$\Gamma_{104}$	$\Lambda_c^- \Lambda_c^+$	< 8.0 $\times 10^{-5}$	CL=95%
$\Gamma_{105}$	$\gamma \gamma$	< 3.1 $\times 10^{-6}$	CL=90%
$\Gamma_{106}$	$\phi \gamma$	( 3.52 $\pm$ 0.34 ) $\times 10^{-5}$	

Lepton Family number (LF) violating modes or  $\Delta B = 1$  weak neutral current (B1) modes

$\Gamma_{107}$	$\mu^+ \mu^-$	B1	( 2.9 $\pm$ 0.7 ) $\times 10^{-9}$	
$\Gamma_{108}$	$e^+ e^-$	B1	< 2.8 $\times 10^{-7}$	CL=90%
$\Gamma_{109}$	$\mu^+ \mu^- \mu^+ \mu^-$	B1	< 1.2 $\times 10^{-8}$	CL=90%
$\Gamma_{110}$	$S P, S \rightarrow \mu^+ \mu^-, P \rightarrow \mu^+ \mu^-$	B1 [b]	< 1.2 $\times 10^{-8}$	CL=90%



$\Gamma_{111}$	$\phi(1020)\mu^+\mu^-$	$B1$	$(8.2 \pm 1.2) \times 10^{-7}$	
$\Gamma_{112}$	$\pi^+\pi^-\mu^+\mu^-$	$B1$	$(8.4 \pm 1.7) \times 10^{-8}$	
$\Gamma_{113}$	$\phi\nu\bar{\nu}$	$B1$	$< 5.4 \times 10^{-3}$	CL=90%
$\Gamma_{114}$	$e^\pm\mu^\mp$	$LF$	$[c] < 1.1 \times 10^{-8}$	CL=90%

- [a] Not a pure measurement. See note at head of  $B_s^0$  Decay Modes.  
 [b] Here  $S$  and  $P$  are the hypothetical scalar and pseudoscalar particles with masses of 2.5 GeV/ $c^2$  and 214.3 MeV/ $c^2$ , respectively.  
 [c] The value is for the sum of the charge states or particle/antiparticle states indicated.

### CONSTRAINED FIT INFORMATION

An overall fit to 10 branching ratios uses 16 measurements and one constraint to determine 7 parameters. The overall fit has a  $\chi^2 = 3.3$  for 10 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_{12}$	28				
$x_{14}$	92	25			
$x_{44}$	0	0	0		
$x_{63}$	0	0	0	62	
$x_{89}$	0	0	0	28	17
	$x_{10}$	$x_{12}$	$x_{14}$	$x_{44}$	$x_{53}$

### $B_s^0$ BRANCHING RATIOS

$\Gamma(D_s^- \text{ anything}) / \Gamma_{\text{total}}$				$\Gamma_1 / \Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.93 ± 0.25 OUR AVERAGE</b>				
0.91 ± 0.18 ± 0.41		1 DRUTSKOY	07 BELL	$e^+e^- \rightarrow \gamma(4S)$
0.81 ± 0.24 ± 0.22	90	2 BUSKULIC	96E ALEP	$e^+e^- \rightarrow Z$
1.56 ± 0.58 ± 0.44	147	3 ACTON	92N OPAL	$e^+e^- \rightarrow Z$

- <sup>1</sup> The extraction of this result takes into account the correlation between the measurements of  $B(\gamma(5S) \rightarrow D_s X)$  and  $B(\gamma(5S) \rightarrow D^0 X)$ .  
<sup>2</sup> BUSKULIC 96E separate  $c\bar{c}$  and  $b\bar{b}$  sources of  $D_s^+$  mesons using a lifetime tag, subtract generic  $\bar{D} \rightarrow W^+ \rightarrow D_s^+ \text{ events}$ , and obtain  $B(\bar{D} \rightarrow B_s^0) \times B(B_s^0 \rightarrow D_s^- \text{ anything}) = 0.088 \pm 0.020 \pm 0.020$  assuming  $B(D_s \rightarrow \phi\pi) = (3.5 \pm 0.4) \times 10^{-2}$  and PDG 1994 values for the relative partial widths to other  $D_s$  channels. We evaluate using our current values  $B(\bar{D} \rightarrow B_s^0) = 0.107 \pm 0.014$  and  $B(D_s \rightarrow \phi\pi) = 0.036 \pm 0.009$ . Our first error is their experiment's and our second error is that due to  $B(\bar{D} \rightarrow B_s^0)$  and  $B(D_s \rightarrow \phi\pi)$ .  
<sup>3</sup> ACTON 92N assume that excess of  $147 \pm 48 D_s^0$  events over that expected from  $B^0, B^+,$  and  $c\bar{c}$  is all from  $B_s^0$  decay. The product branching fraction is measured to be  $B(\bar{D} \rightarrow B_s^0)B(B_s^0 \rightarrow D_s^- \text{ anything}) \times B(D_s^- \rightarrow \phi\pi^-) = (5.9 \pm 1.9 \pm 1.1) \times 10^{-3}$ . We evaluate using our current values  $B(\bar{D} \rightarrow B_s^0) = 0.107 \pm 0.014$  and  $B(D_s \rightarrow \phi\pi) = 0.036 \pm 0.009$ . Our first error is their experiment's and our second error is that due to  $B(\bar{D} \rightarrow B_s^0)$  and  $B(D_s \rightarrow \phi\pi)$ .

$\Gamma(\ell\nu_e X) / \Gamma_{\text{total}}$				$\Gamma_2 / \Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>9.6 ± 0.8 OUR AVERAGE</b>				
9.6 ± 0.4 ± 0.7	1 OSWALD	13 BELL	$e^+e^- \rightarrow \gamma(5S)$	
9.5 ± 2.5 ± 1.1 −2.0 −1.9	2 LEES	12A BABR	$e^+e^-$	

- <sup>1</sup> The measurement corresponds to the average of the electron and muon branching fractions.  
<sup>2</sup> The measurement corresponds to a branching fraction where the lepton originates from bottom decay and is the average between the electron and muon branching fractions. LEES 12A uses the correlation of the production of  $\phi$  mesons in association with a lepton in  $e^+e^-$  data taken at center-of-mass energies between 10.54 and 11.2 GeV.

$\Gamma(e^+\nu X^-) / \Gamma_{\text{total}}$				$\Gamma_3 / \Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>9.1 ± 0.5 ± 0.6</b>	OSWALD	13 BELL	$e^+e^- \rightarrow \gamma(5S)$	

$\Gamma(\mu^+\nu X^-) / \Gamma_{\text{total}}$				$\Gamma_4 / \Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>10.2 ± 0.6 ± 0.8</b>	OSWALD	13 BELL	$e^+e^- \rightarrow \gamma(5S)$	

$\Gamma(D_s^- \ell^+ \nu_\ell \text{ anything}) / \Gamma_{\text{total}}$				$\Gamma_5 / \Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.1 ± 1.3 OUR AVERAGE</b>				
8.2 ± 0.2 ± 1.5		1 OSWALD	15 BELL	$e^+e^- \rightarrow \gamma(5S)$
7.6 ± 1.2 ± 2.1	134	2 BUSKULIC	95O ALEP	$e^+e^- \rightarrow Z$
10.7 ± 4.3 ± 2.9		3 ABREU	92M DLPH	$e^+e^- \rightarrow Z$
10.3 ± 3.6 ± 2.8	18	4 ACTON	92N OPAL	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13 ± 4 ± 4	27	5 BUSKULIC	92E ALEP	$e^+e^- \rightarrow Z$

- The values and averages in this section serve only to show what values result if one assumes our  $B(\bar{D} \rightarrow B_s^0)$ . They cannot be thought of as measurements since the underlying product branching fractions were also used to determine  $B(\bar{D} \rightarrow B_s^0)$  as described in the note on "Production and Decay of  $b$ -Flavored Hadrons."
- <sup>1</sup> Obtains  $B_s \rightarrow D_s X e \nu$ , and  $D_s X \mu \nu$  separately, then combines them by assuming systematic uncertainties are fully correlated, except for the one on lepton identification. The third uncertainty adds in quadrature systematic uncertainties from external sources (number of  $B_s$  events, and  $D_s^{(*)}$  branching fractions). OSWALD 15 also measures the cross-section  $\sigma(e^+e^- \rightarrow B_s^{(*)} \bar{B}_s^{(*)}) = 53.8 \pm 1.4 \pm 5.3$  pb at  $\sqrt{s} = 10.86$  GeV.  
<sup>2</sup> BUSKULIC 95O use  $D_s \ell$  correlations. The measured product branching ratio is  $B(\bar{D} \rightarrow B_s) \times B(B_s \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything}) = (0.82 \pm 0.09 \pm 0.13 \pm 0.14)\%$  assuming  $B(D_s \rightarrow \phi\pi) = (3.5 \pm 0.4) \times 10^{-2}$  and PDG 1994 values for the relative partial widths to the six other  $D_s$  channels used in this analysis. Combined with results from  $\gamma(4S)$  experiments this can be used to extract  $B(\bar{D} \rightarrow B_s) = (11.0 \pm 1.2 \pm 2.5 \pm 2.6)\%$ . We evaluate using our current values  $B(\bar{D} \rightarrow B_s^0) = 0.107 \pm 0.014$  and  $B(D_s \rightarrow \phi\pi) = 0.036 \pm 0.009$ . Our first error is their experiment's and our second error is that due to  $B(\bar{D} \rightarrow B_s^0)$  and  $B(D_s \rightarrow \phi\pi)$ .  
<sup>3</sup> ABREU 92M measured muons only and obtained product branching ratio  $B(Z \rightarrow b \text{ or } \bar{b}) \times B(\bar{D} \rightarrow B_s) \times B(B_s \rightarrow D_s \mu^+ \nu_\mu \text{ anything}) \times B(D_s \rightarrow \phi\pi) = (18 \pm 8) \times 10^{-5}$ . We evaluate using our current values  $B(\bar{D} \rightarrow B_s^0) = 0.107 \pm 0.014$  and  $B(D_s \rightarrow \phi\pi) = 0.036 \pm 0.009$ . Our first error is their experiment's and our second error is that due to  $B(\bar{D} \rightarrow B_s^0)$  and  $B(D_s \rightarrow \phi\pi)$ .  
<sup>4</sup> ACTON 92N is measured using  $D_s \rightarrow \phi\pi^+$  and  $K^*(892)^0 K^+$  events. The product branching fraction measured is measured to be  $B(\bar{D} \rightarrow B_s^0)B(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything}) \times B(D_s^- \rightarrow \phi\pi^-) = (3.9 \pm 1.1 \pm 0.8) \times 10^{-4}$ . We evaluate using our current values  $B(\bar{D} \rightarrow B_s^0) = 0.107 \pm 0.014$  and  $B(D_s \rightarrow \phi\pi) = 0.036 \pm 0.009$ . Our first error is their experiment's and our second error is that due to  $B(\bar{D} \rightarrow B_s^0)$  and  $B(D_s \rightarrow \phi\pi)$ .  
<sup>5</sup> BUSKULIC 92E is measured using  $D_s \rightarrow \phi\pi^+$  and  $K^*(892)^0 K^+$  events. They use  $2.7 \pm 0.7\%$  for the  $\phi\pi^+$  branching fraction. The average product branching fraction is measured to be  $B(\bar{D} \rightarrow B_s^0)B(B_s^0 \rightarrow D_s^- \ell^+ \nu_\ell \text{ anything}) = 0.020 \pm 0.0055 \pm 0.005 - 0.006$ . We evaluate using our current values  $B(\bar{D} \rightarrow B_s^0) = 0.107 \pm 0.014$  and  $B(D_s \rightarrow \phi\pi) = 0.036 \pm 0.009$ . Our first error is their experiment's and our second error is that due to  $B(\bar{D} \rightarrow B_s^0)$  and  $B(D_s \rightarrow \phi\pi)$ . Superseded by BUSKULIC 95O.

$\Gamma(D_s^- \ell^+ \nu_\ell \text{ anything}) / \Gamma_{\text{total}}$				$\Gamma_6 / \Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>5.4 ± 0.4 ± 1.0</b>	1 OSWALD	15 BELL	$e^+e^- \rightarrow \gamma(5S)$	

- <sup>1</sup> Obtains  $B_s \rightarrow D_s^* X e \nu$ , and  $D_s^* X \mu \nu$  separately, then combines them by assuming systematic uncertainties are fully correlated, except for the one on lepton identification. The third uncertainty adds in quadrature systematic uncertainties from external sources (number of  $B_s$  events, and  $D_s^{(*)}$  branching fractions). OSWALD 15 also measures the cross-section  $\sigma(e^+e^- \rightarrow B_s^{(*)} \bar{B}_s^{(*)}) = 53.8 \pm 1.4 \pm 5.3$  pb at  $\sqrt{s} = 10.86$  GeV.

$\Gamma(D_{s1}(2536)^- \mu^+ \nu_\mu, D_{s1}^- \rightarrow D^{*-} K_S^0) / \Gamma_{\text{total}}$				$\Gamma_7 / \Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b>2.6 ± 0.7 ± 0.1</b>	1 ABAZOV	09G D0	$p\bar{p}$ at 1.96 TeV	

- <sup>1</sup> ABAZOV 09G reports  $[\Gamma(B_s^0 \rightarrow D_{s1}(2536)^- \mu^+ \nu_\mu, D_{s1}^- \rightarrow D^{*-} K_S^0) / \Gamma_{\text{total}}] \times [B(\bar{D} \rightarrow B_s^0)] = (2.66 \pm 0.52 \pm 0.45) \times 10^{-4}$  which we divide by our best value  $B(\bar{D} \rightarrow B_s^0) = (10.3 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D_{s1}(2536)^- X \mu^+ \nu, D_{s1}^- \rightarrow \bar{D}^0 K^+) / \Gamma(D_s^- \ell^+ \nu_\ell \text{ anything})$				$\Gamma_8 / \Gamma_5$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>5.4 ± 1.2 ± 0.5</b>	AAIJ	11A LHCB	$pp$ at 7 TeV	

$\Gamma(D_{s2}(2573)^- X \mu^+ \nu, D_{s2}^- \rightarrow \bar{D}^0 K^+) / \Gamma(D_s^- \ell^+ \nu_\ell \text{ anything})$				$\Gamma_9 / \Gamma_5$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>3.3 ± 1.0 ± 0.4</b>	AAIJ	11A LHCB	$pp$ at 7 TeV	

$\Gamma(D_{s1}(2536)^- X \mu^+ \nu, D_{s1}^- \rightarrow \bar{D}^0 K^+) / \Gamma(D_{s2}(2573)^- X \mu^+ \nu, D_{s2}^- \rightarrow \bar{D}^0 K^+)$				$\Gamma_8 / \Gamma_9$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.61 ± 0.14 ± 0.05	1 AAIJ	11A LHCB	$pp$ at 7 TeV	

<sup>1</sup> Not independent of other AAIJ 11A measurements.





# Meson Particle Listings

$B_S^0$

$\Gamma(D_{s1}^*(2860)^-\pi^+, D_{s1}^* \rightarrow \bar{D}^0 K^-)/\Gamma_{\text{total}}$	$\Gamma_{38}/\Gamma$		
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.0 \pm 1.2 \pm 3.4</math></b>	<sup>1</sup> AAIJ	14BH LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Uses Dalitz plot analysis of  $B_S^0 \rightarrow \bar{D}^0 K^- \pi^+$  decays.

$\Gamma(D_{s3}^*(2860)^-\pi^+, D_{s3}^* \rightarrow \bar{D}^0 K^-)/\Gamma_{\text{total}}$	$\Gamma_{39}/\Gamma$		
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.2 \pm 0.1 \pm 0.6</math></b>	<sup>1</sup> AAIJ	14BH LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Uses Dalitz plot analysis of  $B_S^0 \rightarrow \bar{D}^0 K^- \pi^+$  decays.

$\Gamma(\bar{D}^0 K^+ K^-)/\Gamma_{\text{total}}$	$\Gamma_{40}/\Gamma$		
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>4.4 \pm 1.7 \pm 1.1</math></b>	<sup>1,2</sup> AAIJ	12AMLHCB	$pp$ at 7 TeV

<sup>1</sup> AAIJ 12AM reports  $[\Gamma(B_S^0 \rightarrow \bar{D}^0 K^+ K^-)/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow \bar{D}^0 K^+ K^-)] = 0.90 \pm 0.27 \pm 0.20$  which we multiply by our best value  $\text{B}(B^0 \rightarrow \bar{D}^0 K^+ K^-) = (4.9 \pm 1.2) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Uses  $\text{B}(b \rightarrow B_S^0)/\text{B}(b \rightarrow B^0) = 0.267^{+0.023}_{-0.020}$  measured by the same authors.

$\Gamma(\bar{D}^0 f_0(980))/\Gamma_{\text{total}}$		$\Gamma_{41}/\Gamma$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-6}$	90	AAIJ	15AG LHCb	$pp$ at 7, 8 TeV

$\Gamma(\bar{D}^0 \phi)/\Gamma(\bar{D}^0 K^*(892)^0)$	$\Gamma_{42}/\Gamma_{27}$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.069 \pm 0.013 \pm 0.007</math></b>	AAIJ	13BX LHCb	$pp$ at 7 TeV

$\Gamma(D^{*\mp}\pi^\pm)/\Gamma_{\text{total}}$					$\Gamma_{43}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6.1 \times 10^{-6}</math></b>	90	<sup>1</sup> AAIJ	13AL LHCb	$pp$ at 7 TeV	

<sup>1</sup> Uses  $f_s/f_d = 0.256 \pm 0.020$  and  $\text{B}(B^0 \rightarrow D^{*-} \pi^+) = (2.76 \pm 0.13) \times 10^{-3}$ .

$\Gamma(J/\psi(1S)\phi)/\Gamma_{\text{total}}$					$\Gamma_{44}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.07 ± 0.08 OUR FIT</b>					
<b>1.10 ± 0.09 OUR AVERAGE</b>					

<sup>1</sup> 0.05  $0 \pm 0.013 \pm 0.104$   
<sup>2</sup> 1.25  $\pm 0.07 \pm 0.23$   
<sup>3</sup> 1.4  $\pm 0.5 \pm 0.1$   
• • • We do not use the following data for averages, fits, limits, etc. • • •

<6	1	<sup>4</sup> AKERS	94J OPAL	$e^+e^- \rightarrow Z$
seen	14	<sup>5</sup> ABE	93F CDF	$p\bar{p}$ at 1.8 TeV
seen	1	<sup>6</sup> ACTON	92N OPAL	Sup. by AKERS 94J

<sup>1</sup> Uses  $f_s/f_d = 0.256 \pm 0.020$  and  $\text{B}(B^+ \rightarrow J/\psi K^+) = (10.18 \pm 0.42) \times 10^{-4}$ .  
<sup>2</sup> Uses  $f_s = (17.2 \pm 3.0)\%$  as the fraction of  $\Upsilon(5S)$  decaying to  $B_S^{(*)}\bar{B}_S^{(*)}$ .  
<sup>3</sup> ABE 96Q reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)\phi)/\Gamma_{\text{total}}] \times [\Gamma(\bar{B} \rightarrow B_S^0)/[\Gamma(\bar{B} \rightarrow B^+) + \Gamma(\bar{B} \rightarrow B^0)]] = (0.185 \pm 0.055 \pm 0.020) \times 10^{-3}$  which we divide by our best value  $\Gamma(\bar{B} \rightarrow B_S^0)/[\Gamma(\bar{B} \rightarrow B^+) + \Gamma(\bar{B} \rightarrow B^0)] = 0.130 \pm 0.008$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>4</sup> AKERS 94J sees one event and measures the limit on the product branching fraction  $r(\bar{B} \rightarrow B_S^0) \cdot \text{B}(B_S^0 \rightarrow J/\psi(1S)\phi) < 7 \times 10^{-4}$  at CL = 90%. We divide by  $\text{B}(\bar{B} \rightarrow B_S^0) = 0.112$ .  
<sup>5</sup> ABE 93F measured using  $J/\psi(1S) \rightarrow \mu^+\mu^-$  and  $\phi \rightarrow K^+K^-$ .  
<sup>6</sup> In ACTON 92N a limit on the product branching fraction is measured to be  $r(\bar{B} \rightarrow B_S^0) \cdot \text{B}(B_S^0 \rightarrow J/\psi(1S)\phi) \leq 0.22 \times 10^{-2}$ .

$\Gamma(J/\psi(1S)\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{45}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN
$<1.2 \times 10^{-3}$	90	<sup>1</sup> ACCIARRI	97C L3

<sup>1</sup> ACCIARRI 97C assumes  $B^0$  production fraction (39.5  $\pm$  4.0%) and  $B_S$  (12.0  $\pm$  3.0%).

$\Gamma(J/\psi(1S)\eta)/\Gamma_{\text{total}}$		$\Gamma_{46}/\Gamma$		
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>3.9 \pm 0.7</math> OUR AVERAGE</b>		Error includes scale factor of 1.4.		
$3.6 \pm 0.5$ LQ		1. AAL	12	USP
				1.7 T.M.

<sup>1</sup> AAIJ 13A reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)\eta)/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow J/\psi(1S)\rho^0)] = 14.0 \pm 3.6 \pm 0.5 \pm 0.2$   
<sup>2</sup> LI 12 BELL  $e^+e^- \rightarrow \Upsilon(4S)$   
<sup>3</sup> ACCIARRI 97C L3

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<38 90  
<sup>1</sup> AAIJ 13A reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)\eta)/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow J/\psi(1S)\rho^0)] = 14.0 \pm 1.2^{+1.1+1.1}_{-1.5-1.0}$  which we multiply by our best value  $\text{B}(B^0 \rightarrow J/\psi(1S)\rho^0) = (2.54 \pm 0.14) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Observed for the first time with significances over 10  $\sigma$ . The second error are total systematic uncertainties including the error on  $\text{N}(B_S^{(*)}\bar{B}_S^{(*)})$ .  
<sup>3</sup> ACCIARRI 97C assumes  $B^0$  production fraction (39.5  $\pm$  4.0%) and  $B_S$  (12.0  $\pm$  3.0%).

$\Gamma(J/\psi(1S) K_S^0)/\Gamma_{\text{total}}$	$\Gamma_{47}/\Gamma$		
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.89 \pm 0.12</math> OUR AVERAGE</b>			

<sup>1</sup> 1.88  $\pm 0.14 \pm 0.07$   
<sup>2</sup> 1.91  $\pm 0.15 \pm 0.13$   
<sup>3</sup> 1.9  $\pm 0.4 \pm 0.1$   
• • • We do not use the following data for averages, fits, limits, etc. • • •

1.91 $\pm 0.25$ $-0.24 \pm 0.13$	<sup>4</sup> AAIJ	12o LHCb	Repl. by AAIJ 13AB
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<sup>1</sup> AAIJ 15AL reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)K_S^0)/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow J/\psi(1S)K_S^0)] = (4.31 \pm 0.17 \pm 0.12 \pm 0.25) \times 10^{-2}$  which we multiply by our best value  $\text{B}(B^0 \rightarrow J/\psi(1S)K_S^0) = (4.36 \pm 0.16) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> AAIJ 13AB reports  $(1.97 \pm 0.14 \pm 0.07 \pm 0.15 \pm 0.08) \times 10^{-5}$  from a measurement of  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)K_S^0)/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow J/\psi(1S)K^0)] \times [\Gamma(\bar{B} \rightarrow B_S^0)/\Gamma(\bar{B} \rightarrow B^0)]$  assuming  $\text{B}(B^0 \rightarrow J/\psi(1S)K^0) = (8.98 \pm 0.35) \times 10^{-4}$ ,  $\Gamma(\bar{B} \rightarrow B_S^0)/\Gamma(\bar{B} \rightarrow B^0) = 0.256 \pm 0.020$ , which we rescale to our best values  $\text{B}(B^0 \rightarrow J/\psi(1S)K^0) = (8.73 \pm 0.32) \times 10^{-4}$ ,  $\Gamma(\bar{B} \rightarrow B_S^0)/\Gamma(\bar{B} \rightarrow B^0) = 0.256 \pm 0.014$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.  
<sup>3</sup> AALTONEN 11A reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)K_S^0)/\Gamma_{\text{total}}] \times [\text{B}(\bar{B} \rightarrow B_S^0)] / [\text{B}(\bar{B} \rightarrow B^0)] / [\text{B}(B^0 \rightarrow J/\psi(1S)K_S^0)] = (1.09 \pm 0.19 \pm 0.11) \times 10^{-2}$  which we multiply or divide by our best values  $\text{B}(\bar{B} \rightarrow B_S^0) = (10.3 \pm 0.5) \times 10^{-2}$ ,  $\text{B}(\bar{B} \rightarrow B^0) = (40.4 \pm 0.6) \times 10^{-2}$ ,  $\text{B}(B^0 \rightarrow J/\psi(1S)K_S^0) = 1/2 \times \text{B}(B^0 \rightarrow J/\psi(1S)K^0) = 1/2 \times (8.73 \pm 0.32) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.  
<sup>4</sup> AAIJ 12o reports  $(1.83 \pm 0.21 \pm 0.10 \pm 0.14 \pm 0.07) \times 10^{-5}$  from a measurement of  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)K_S^0)/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow J/\psi(1S)K^0)] \times [\Gamma(\bar{B} \rightarrow B_S^0)/\Gamma(\bar{B} \rightarrow B^0)]$  assuming  $\text{B}(B^0 \rightarrow J/\psi(1S)K^0) = (8.71 \pm 0.32) \times 10^{-4}$ ,  $\Gamma(\bar{B} \rightarrow B_S^0)/\Gamma(\bar{B} \rightarrow B^0) = 0.267^{+0.021}_{-0.02}$ , which we rescale to our best values  $\text{B}(B^0 \rightarrow J/\psi(1S)K^0) = (8.73 \pm 0.32) \times 10^{-4}$ ,  $\Gamma(\bar{B} \rightarrow B_S^0)/\Gamma(\bar{B} \rightarrow B^0) = 0.256 \pm 0.014$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

$\Gamma(J/\psi(1S) K^*(892)^0)/\Gamma_{\text{total}}$	$\Gamma_{48}/\Gamma$		
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>4.14 \pm 0.18 \pm 0.35</math></b>	<sup>1</sup> AAIJ	15AV LHCb	$pp$ at 7, 8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.4 $\pm 0.5$ $-0.4 \pm 0.8$	<sup>2</sup> AAIJ	12AP LHCb	Repl. by AAIJ 15AV
8 $\pm 4 \pm 1$	<sup>3</sup> AALTONEN	11A CDF	$p\bar{p}$ at 1.96 TeV

<sup>1</sup> AAIJ 15AV result combines two measurements with different normalizing modes of  $B^0 \rightarrow J/\psi K^*(892)^0$  and  $B_S^0 \rightarrow J/\psi \phi$ .  
<sup>2</sup> AAIJ 12AP reports  $\text{B}(B_S^0 \rightarrow J/\psi(1S)\bar{K}^*(892)^0)/\text{B}(B^0 \rightarrow J/\psi(1S)K^*(892)^0) = (3.43^{+0.34}_{-0.36} \pm 0.50) \times 10^{-2}$  and  $\text{B}(B^0 \rightarrow J/\psi(1S)K^*(892)^0) = (1.29 \pm 0.05 \pm 0.13) \times 10^{-3}$  after correcting for the contribution from  $K\pi$  S-wave beneath the  $K^*$  peak.  
<sup>3</sup> AALTONEN 11A reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)\bar{K}^*(892)^0)/\Gamma_{\text{total}}] \times [\text{B}(\bar{B} \rightarrow B_S^0)] / [\text{B}(\bar{B} \rightarrow B^0)] / [\text{B}(B^0 \rightarrow J/\psi(1S)K^*(892)^0)] = 0.0168 \pm 0.0024 \pm 0.0068$  which we multiply or divide by our best values  $\text{B}(\bar{B} \rightarrow B_S^0) = (10.3 \pm 0.5) \times 10^{-2}$ ,  $\text{B}(\bar{B} \rightarrow B^0) = (40.4 \pm 0.6) \times 10^{-2}$ ,  $\text{B}(B^0 \rightarrow J/\psi(1S)K^*(892)^0) = (1.28 \pm 0.05) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

$\Gamma(J/\psi(1S)\eta')/\Gamma_{\text{total}}$	$\Gamma_{49}/\Gamma$		
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.3 \pm 0.4</math> OUR AVERAGE</b>			

3.2 $\pm 0.4$ $-0.5 \pm 0.2$	<sup>1</sup> AAIJ	13A LHCb	$pp$ at 7 TeV
3.71 $\pm 0.61 \pm 0.85$ $-0.60$	<sup>2</sup> LI	12 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> AAIJ 13A reports  $[\Gamma(B_S^0 \rightarrow J/\psi(1S)\eta')/\Gamma_{\text{total}}] / [\text{B}(B^0 \rightarrow J/\psi(1S)\rho^0)] = 12.7 \pm 1.1^{+0.5+1.0}_{-1.3-0.9}$  which we multiply by our best value  $\text{B}(B^0 \rightarrow J/\psi(1S)\rho^0) = (2.54 \pm 0.14) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Observed for the first time with significances over 10  $\sigma$ . The second error are total systematic uncertainties including the error on  $\text{N}(B_S^{(*)}\bar{B}_S^{(*)})$ .

$\Gamma(J/\psi(1S)\eta')/\Gamma(J/\psi(1S)\eta)$	$\Gamma_{49}/\Gamma_{46}$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.87 \pm 0.06</math> OUR AVERAGE</b>			
$0.902 \pm 0.072 \pm 0.045$	<sup>1</sup> AAI1	15D 1 HCB	pp at 7.8 TeV

0.902 $\pm 0.072 \pm 0.045$	<sup>1</sup> AAIJ	15D LHCb	$pp$ at 7, 8 TeV
0.90 $\pm 0.09 \pm 0.06$ $-0.02$	<sup>2</sup> AAIJ	13A LHCb	$pp$ at 7 TeV
0.73 $\pm 0.14 \pm 0.02$	<sup>2</sup> LI	12 BELL	$e^+e^- \rightarrow \Upsilon(4S)$

<sup>1</sup> Uses  $J/\psi \rightarrow \mu^+\mu^-$ ,  $\eta' \rightarrow \rho^0\gamma$ , and  $\eta' \rightarrow \eta\pi^+\pi^-$  decays.  
<sup>2</sup> Strongly correlated with measurements of  $\Gamma(J/\psi(1S)\eta)/\Gamma$  and  $\Gamma(J/\psi(1S)\eta')/\Gamma$  reported in the same reference.













# Meson Particle Listings

$B_s^0$

$\phi_{\parallel}$ in $B_s^0 \rightarrow J/\psi(1S)\phi$				
VALUE (rad)	DOCUMENT ID	TECN	COMMENT	
<b><math>3.23^{+0.10}_{-0.14}</math> OUR AVERAGE</b>				
$3.26^{+0.10+0.06}_{-0.17-0.07}$	AAIJ	15i	LHCB	$pp$ at 7, 8 TeV
$3.15 \pm 0.22$	<sup>1</sup> ABAZOV	12d	D0	$p\overline{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.72^{+1.12}_{-0.27} \pm 0.26$	ABAZOV	09E	D0	Repl. by ABAZOV 12d
<sup>1</sup> The error includes both statistical and systematic uncertainties.				

$\phi_{\perp}$ in $B_s^0 \rightarrow J/\psi(1S)\phi$				
VALUE (rad)	DOCUMENT ID	TECN	COMMENT	
<b><math>3.16 \pm 0.24</math> OUR AVERAGE</b>	Error includes scale factor of 1.6.			
$3.08^{+0.14}_{-0.15} \pm 0.06$	AAIJ	15i	LHCB	$pp$ at 7, 8 TeV
$3.89 \pm 0.47 \pm 0.11$	<sup>1</sup> AAD	14u	ATLS	$pp$ at 7 TeV
<sup>1</sup> Measured using a tagged, time-dependent angular analysis of $B_s^0 \rightarrow J/\psi\phi$ decays.				

$\Gamma_L/\Gamma$ for $B_s^0 \rightarrow J/\psi(1S)\overline{K}^*(892)^0$				
Longitudinal polarization fraction, equals to $f_L$ using notation of “Polarization in $B$ decays” review.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.497 \pm 0.025 \pm 0.025</math></b>	AAIJ	15AV	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.50 \pm 0.08 \pm 0.02$	<sup>1</sup> AAIJ	12AP	LHCB	Repl. by AAIJ 15AV
<sup>1</sup> The non-resonant $K\pi$ background contributions are subtracted. Also reports an S-wave amplitude $ A_S ^2 = 0.07^{+0.15}_{-0.07}$ .				

$\Gamma_{\parallel}/\Gamma$ for $B_s^0 \rightarrow J/\psi(1S)\overline{K}^*(892)^0$				
Parallel polarization fraction, equals to $1 - f_L - f_{\perp}$ using notation of “Polarization in $B$ decays” review.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.179 \pm 0.027 \pm 0.013</math></b>	AAIJ	15AV	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.19^{+0.10}_{-0.08} \pm 0.02$	<sup>1</sup> AAIJ	12AP	LHCB	Repl. by AAIJ 15AV
<sup>1</sup> The non-resonant $K\pi$ background contributions are subtracted. Also reports an S-wave amplitude $ A_S ^2 = 0.07^{+0.15}_{-0.07}$ .				

$\Gamma_{\parallel}/\Gamma$ of $K^*(892)^0$ in $B_s^0 \rightarrow \psi(2S)\overline{K}^*(892)^0$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.524 \pm 0.056 \pm 0.029</math></b>	AAIJ	15u	LHCB	$pp$ at 7, 8 TeV

$\Gamma_L/\Gamma$ in $B_s^0 \rightarrow \phi\phi$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.362 \pm 0.014</math> OUR AVERAGE</b>				
$0.364 \pm 0.012 \pm 0.009$	AAIJ	14AE	LHCB	$pp$ at 7, 8 TeV
$0.348 \pm 0.041 \pm 0.021$	AALTONEN	11AN	CDF	$p\overline{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.365 \pm 0.022 \pm 0.012$	AAIJ	12P	LHCB	Repl. by AAIJ 14AE

$\Gamma_{\perp}/\Gamma$ in $B_s^0 \rightarrow \phi\phi$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.309 \pm 0.015</math> OUR AVERAGE</b>	Error includes scale factor of 1.1.			
$0.305 \pm 0.013 \pm 0.005$	AAIJ	14AE	LHCB	$pp$ at 7, 8 TeV
$0.365 \pm 0.044 \pm 0.027$	AALTONEN	11AN	CDF	$p\overline{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.291 \pm 0.024 \pm 0.010$	AAIJ	12P	LHCB	Repl. by AAIJ 14AE

$\phi_{\parallel}$ in $B_s^0 \rightarrow \phi\phi$				
VALUE (rad)	DOCUMENT ID	TECN	COMMENT	
<b><math>2.55 \pm 0.11</math> OUR AVERAGE</b>				
$2.54 \pm 0.07 \pm 0.09$	<sup>1</sup> AAIJ	14AE	LHCB	$pp$ at 7, 8 TeV
$2.71^{+0.31}_{-0.36} \pm 0.22$	<sup>2</sup> AALTONEN	11AN	CDF	$p\overline{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2.57 \pm 0.15 \pm 0.06$	<sup>3</sup> AAIJ	12P	LHCB	Repl. by AAIJ 14AE
<sup>1</sup> AAIJ 14AE reports measurement of $\phi_{\perp}$ and $\phi_{\perp} - \phi_{\parallel}$ , which we convert into $\phi_{\parallel}$ . Statistical uncertainty includes correlation between measured parameters, while systematic uncertainties are assumed uncorrelated.				
<sup>2</sup> AALTONEN 11AN quotes $\cos\phi_{\parallel} = -0.91^{+0.15}_{-0.13} \pm 0.09$ which we convert to $\phi_{\parallel}$ taking the smaller solution.				
<sup>3</sup> AAIJ 12P quotes $\cos\phi_{\parallel} = -0.844 \pm 0.068 \pm 0.029$ which we convert to $\phi_{\parallel}$ , taking the smaller solution.				

$\phi_{\perp}$ in $B_s^0 \rightarrow \phi\phi$				
VALUE (rad)	DOCUMENT ID	TECN	COMMENT	
<b><math>2.67 \pm 0.23 \pm 0.07</math></b>	AAIJ	14AE	LHCB	$pp$ at 7, 8 TeV

$\Gamma_L/\Gamma$ in $B_s^0 \rightarrow K^{*0}\overline{K}^{*0}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.201 \pm 0.057 \pm 0.040</math></b>	<sup>1</sup> AAIJ	15AF	LHCB	$pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.31 \pm 0.12 \pm 0.04$	AAIJ	12F	LHCB	Repl. by AAIJ 15AF
<sup>1</sup> Measured in angular analysis, which takes into account S-wave contributions.				

$\Gamma_L/\Gamma$ in $B_s^0 \rightarrow K^{*0}\overline{K}^{*0}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.38 \pm 0.11 \pm 0.04</math></b>	AAIJ	12F	LHCB	$pp$ at 7 TeV

$\Gamma_{\parallel}/\Gamma$ in $B_s^0 \rightarrow K^*(892)^0\overline{K}^*(892)^0$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.215 \pm 0.046 \pm 0.015</math></b>	AAIJ	15AF	LHCB	$pp$ at 7 TeV

$\phi_{\parallel}$ in $B_s^0 \rightarrow K^*(892)^0\overline{K}^*(892)^0$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>5.31 \pm 0.24 \pm 0.14</math></b>	AAIJ	15AF	LHCB	$pp$ at 7 TeV

$\Gamma_L/\Gamma$ in $B_s^0 \rightarrow \phi\overline{K}^{*0}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.51 \pm 0.15 \pm 0.07</math></b>	AAIJ	13BW	LHCB	$pp$ at 7 TeV

$\Gamma_{\parallel}/\Gamma$ in $B_s^0 \rightarrow \phi\overline{K}^{*0}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.21 \pm 0.11 \pm 0.02</math></b>	AAIJ	13BW	LHCB	$pp$ at 7 TeV

$\phi_{\parallel}$ in $B_s^0 \rightarrow \phi\overline{K}^{*0}$				
VALUE (rad)	DOCUMENT ID	TECN	COMMENT	
<b><math>1.75 \pm 0.53 \pm 0.29</math></b>	<sup>1</sup> AAIJ	13BW	LHCB	$pp$ at 7 TeV
<sup>1</sup> Measures $\cos(\phi_{\parallel}) = -0.18 \pm 0.52 \pm 0.29$ , which we convert to $\phi_{\parallel}$ by taking the smaller solution.				

$F_L(B_s^0 \rightarrow \phi\mu^+\mu^-)$ ( $0.10 < q^2 < 2.00 \text{ GeV}^2/c^4$ )				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.20^{+0.08}_{-0.09} \pm 0.02</math></b>	AAIJ	15AQ	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.37^{+0.19}_{-0.17} \pm 0.07$	AAIJ	13x	LHCB	Repl. by AAIJ 15AQ

$F_L(B_s^0 \rightarrow \phi\mu^+\mu^-)$ ( $2.00 < q^2 < 5.0 \text{ GeV}^2/c^4$ )				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.68^{+0.16}_{-0.13} \pm 0.03</math></b>	AAIJ	15AQ	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.53^{+0.25}_{-0.23} \pm 0.10$	<sup>1</sup> AAIJ	13x	LHCB	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $2.0 < q^2 < 4.3 \text{ GeV}^2/c^4$ .				

$F_L(B_s^0 \rightarrow \phi\mu^+\mu^-)$ ( $5.0 < q^2 < 8.0 \text{ GeV}^2/c^4$ )				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.54^{+0.10}_{-0.09} \pm 0.02</math></b>	AAIJ	15AQ	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.81^{+0.11}_{-0.13} \pm 0.05$	<sup>1</sup> AAIJ	13x	LHCB	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $4.3 < q^2 < 8.68 \text{ GeV}^2/c^4$ .				

$F_L(B_s^0 \rightarrow \phi\mu^+\mu^-)$ ( $11.0 < q^2 < 12.5 \text{ GeV}^2/c^4$ )				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.29 \pm 0.11 \pm 0.04</math></b>	AAIJ	15AQ	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.33^{+0.14}_{-0.12} \pm 0.06$	<sup>1</sup> AAIJ	13x	LHCB	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $10.09 < q^2 < 12.90 \text{ GeV}^2/c^4$ .				

$F_L(B_s^0 \rightarrow \phi\mu^+\mu^-)$ ( $15.0 < q^2 < 17.0 \text{ GeV}^2/c^4$ )				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.23^{+0.09}_{-0.08} \pm 0.02</math></b>	AAIJ	15AQ	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.34^{+0.18}_{-0.17} \pm 0.07$	<sup>1</sup> AAIJ	13x	LHCB	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $14.18 < q^2 < 16 \text{ GeV}^2/c^4$ .				

$F_L(B_s^0 \rightarrow \phi\mu^+\mu^-)$ ( $17.0 < q^2 < 19.0 \text{ GeV}^2/c^4$ )				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.40^{+0.13}_{-0.15} \pm 0.02</math></b>	AAIJ	15AQ	LHCB	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.16^{+0.17}_{-0.10} \pm 0.07$	<sup>1</sup> AAIJ	13x	LHCB	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $16.0 < q^2 < 19.0 \text{ GeV}^2/c^4$ .				

See key on page 601

# Meson Particle Listings

 $B_s^0$ 

## $F_L(B_s^0 \rightarrow \phi \mu^+ \mu^-)$ ( $1.00 < q^2 < 6.00 \text{ GeV}^2/c^4$ )

VALUE	DOCUMENT ID	TECN	COMMENT
$0.63^{+0.09}_{-0.09} \pm 0.03$	AAIJ	15AQ LHCb	$pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.56^{+0.17}_{-0.16} \pm 0.09$	AAIJ	13x LHCb	Repl. by AAIJ 15AQ

## $B_s^0\text{-}\bar{B}_s^0$ MIXING

For a discussion of  $B_s^0\text{-}\bar{B}_s^0$  mixing see the note on “ $B^0\text{-}\bar{B}^0$  Mixing” in the  $B^0$  Particle Listings above.

$\chi_s$  is a measure of the time-integrated  $B_s^0\text{-}\bar{B}_s^0$  mixing probability that produced  $B_s^0(\bar{B}_s^0)$  decays as a  $\bar{B}_s^0(B_s^0)$ . Mixing violates  $\Delta B \neq 2$  rule.

$$\chi_s = \frac{x_s^2}{2(1+x_s^2)}$$

$$x_s = \frac{\Delta m_{B_s^0}}{\Gamma_{B_s^0}} = (m_{B_{sH}^0} - m_{B_{sL}^0}) \tau_{B_s^0},$$

where  $H, L$  stand for heavy and light states of two  $B_s^0$  CP eigenstates and

$$\tau_{B_s^0} = \frac{1}{0.5(\Gamma_{B_{sH}^0} + \Gamma_{B_{sL}^0})}.$$

$$\Delta m_{B_s^0} = m_{B_{sH}^0} - m_{B_{sL}^0}$$

$\Delta m_{B_s^0}$  is a measure of  $2\pi$  times the  $B_s^0\text{-}\bar{B}_s^0$  oscillation frequency in time-dependent mixing experiments.

“OUR EVALUATION” is provided by the Heavy Flavor Averaging Group (HFAG) by taking into account correlations between measurements.

VALUE ( $10^{12} \text{ h s}^{-1}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>17.757<math>\pm</math>0.021 OUR EVALUATION</b>				
<b>17.756<math>\pm</math>0.021 OUR AVERAGE</b>				
$17.711^{+0.055}_{-0.057} \pm 0.011$	1	AAIJ	15I LHCb	$pp$ at 7, 8 TeV
$17.768 \pm 0.023 \pm 0.006$	2	AAIJ	13BI LHCb	$pp$ at 7 TeV
$17.93 \pm 0.22 \pm 0.15$	3	AAIJ	13CF LHCb	$pp$ at 7 TeV
$17.63 \pm 0.11 \pm 0.02$	4	AAIJ	12I LHCb	$pp$ at 7 TeV
$17.77 \pm 0.10 \pm 0.07$	5	ABULENCIA,A 06G	CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17–21	90	6	ABAZOV	06B D0 $p\bar{p}$ at 1.96 TeV
$17.31^{+0.33}_{-0.18} \pm 0.07$	7	ABULENCIA	06Q CDF	Repl. by ABULENCIA,A 06G
> 8.0	95	8	ABDALLAH	04J DLPH $e^+e^- \rightarrow Z^0$
> 4.9	95	9	ABDALLAH	04J DLPH $e^+e^- \rightarrow Z^0$
> 8.5	95	10	ABDALLAH	04J DLPH $e^+e^- \rightarrow Z^0$
> 5.0	95	11	ABDALLAH	03B DLPH $e^+e^- \rightarrow Z$
>10.3	95	12	ABE	03 SLD $e^+e^- \rightarrow Z$
>10.9	95	13	HEISTER	03E ALEP $e^+e^- \rightarrow Z$
> 5.3	95	14	ABE	02V SLD $e^+e^- \rightarrow Z$
> 1.0	95	15	ABBIENDI	01D OPAL $e^+e^- \rightarrow Z$
> 7.4	95	16	ABREU	00Y DLPH Repl. by ABDALLAH 04J
> 4.0	95	17	ABREU,P	00G DLPH $e^+e^- \rightarrow Z$
> 5.2	95	18	ABBIENDI	99S OPAL $e^+e^- \rightarrow Z$
<96	95	19	ABE	99D CDF $p\bar{p}$ at 1.8 TeV
> 5.8	95	20	ABE	99J CDF $p\bar{p}$ at 1.8 TeV
> 9.6	95	21	BARATE	99J ALEP $e^+e^- \rightarrow Z$
> 7.9	95	22	BARATE	98c ALEP Repl. by BARATE 99J
> 3.1	95	23	ACKERSTAFF	97U OPAL Repl. by ABBIENDI 99s
> 2.2	95	24	ACKERSTAFF	97V OPAL Repl. by ABBIENDI 99s
> 6.5	95	25	ADAM	97 DLPH Repl. by ABREU 00Y
> 6.6	95	26	BUSKULIC	96M ALEP Repl. by BARATE 98c
> 2.2	95	24	AKERS	95J OPAL Sup. by ACKERSTAFF 97V
> 5.7	95	27	BUSKULIC	95J ALEP $e^+e^- \rightarrow Z$
> 1.8	95	24	BUSKULIC	94B ALEP $e^+e^- \rightarrow Z$

1 Measured using time-dependent angular analysis of  $B_s^0 \rightarrow J/\psi K^+ K^-$  decays.

2 Measured using  $B_s^0 \rightarrow D_s^- \pi^+$  decays.

3 Measured using  $B_s^0 \rightarrow D_s^- \mu^+ \nu_\mu X$  decays.

4 Measured using  $B_s^0 \rightarrow D_s^- \pi^+$  and  $D_s^- \pi^+ \pi^- \pi^+$  decays.

5 Significance of oscillation signal is  $5.4 \sigma$ . Also reports  $|V_{td}/V_{ts}| = 0.2060 \pm 0.0007^{+0.0081}_{-0.0060}$ .

6 A likelihood scan over the oscillation frequency,  $\Delta m_s$ , gives a most probable value of  $19 \text{ ps}^{-1}$  and a range of  $17 < \Delta m_s < 21 (\text{ps}^{-1})$  at 90% C.L. assuming Gaussian uncertainties. Also excludes  $\Delta m_s < 14.8 \text{ ps}^{-1}$  at 95% C.L.

7 Significance of oscillation signal is 0.2%. Also reported the value  $|V_{td}/V_{ts}| = 0.208^{+0.001+0.008}_{-0.002-0.006}$ .

8 Uses leptons emitted with large momentum transverse to a jet and improved techniques for vertexing and flavor-tagging.

9 Updates of  $D_s$ -lepton analysis.

10 Combined results from all Delphi analyses.

11 Events with a high transverse momentum lepton were removed and an inclusively reconstructed vertex was required.

12 ABE 03 uses the novel “charge dipole” technique to reconstruct separate secondary and tertiary vertices originating from the  $B \rightarrow D$  decay chain. The analysis excludes  $\Delta m_s < 4.9 \text{ ps}^{-1}$  and  $7.9 < \Delta m_s < 10.3 \text{ ps}^{-1}$ .

13 Three analyses based on complementary event selections: (1) fully-reconstructed hadronic decays; (2) semileptonic decays with  $D_s$  exclusively reconstructed; (3) inclusive semileptonic decays.

14 ABE 02v uses exclusively reconstructed  $D_s^-$  mesons and excludes  $\Delta m_s < 1.4 \text{ ps}^{-1}$  and  $2.4 < \Delta m_s < 5.3 \text{ ps}^{-1}$  at 95%CL.

15 Uses fully or partially reconstructed  $D_s \ell$  vertices and a mixing tag as a flavor tagging.

16 Replaced by ABDALLAH 04A. Uses  $D_s^- \ell^+$ , and  $\phi \ell^+$  vertices, and a multi-variable discriminant as a flavor tagging.

17 Uses inclusive  $D_s$  vertices and fully reconstructed  $B_s$  decays and a multi-variable discriminant as a flavor tagging.

18 Uses  $\ell\text{-}Q_{\text{hem}}$  and  $\ell\text{-}\ell$ .

19 ABE 99d assumes  $\tau_{B_s^0} = 1.55 \pm 0.05 \text{ ps}$  and  $\Delta\Gamma/\Delta m = (5.6 \pm 2.6) \times 10^{-3}$ .

20 ABE 99j uses  $\phi \ell\text{-}\ell$  correlation.

21 BARATE 99j uses combination of an inclusive lepton and  $D_s^-$ -based analyses.

22 BARATE 98c combines results from  $D_s \ell\text{-}\ell/Q_{\text{hem}}$ ,  $D_s \ell\text{-}K$  in the same side,  $D_s \ell\text{-}\ell/Q_{\text{hem}}$  and  $D_s \ell\text{-}K$  in the same side.

23 Uses  $\ell\text{-}Q_{\text{hem}}$ .

24 Uses  $\ell\text{-}\ell$ .

25 ADAM 97 combines results from  $D_s \ell\text{-}Q_{\text{hem}}$ ,  $\ell\text{-}Q_{\text{hem}}$ , and  $\ell\text{-}\ell$ .

26 BUSKULIC 96M uses  $D_s$  lepton correlations and lepton, kaon, and jet charge tags.

27 BUSKULIC 95i uses  $\ell\text{-}Q_{\text{hem}}$ . They find  $\Delta m_s > 5.6 [ > 6.1]$  for  $f_s=10\%$  [12%]. We interpolate to our central value  $f_s=10.5\%$ .

$$\chi_s = \Delta m_{B_s^0} / \Gamma_{B_s^0}$$

This is derived by the Heavy Flavor Averaging Group (HFAG) from the results on  $\Delta m_{B_s^0}$  and “OUR EVALUATION” of the  $B_s^0$  mean lifetime.

VALUE	DOCUMENT ID
<b>26.81<math>\pm</math>0.10 OUR EVALUATION</b>	

## $\chi_s$

This is a  $B_s^0\text{-}\bar{B}_s^0$  integrated mixing parameter derived from  $\chi_s$  above and OUR EVALUATION of  $\Delta\Gamma_{B_s^0}/\Gamma_{B_s^0}$ .

VALUE	DOCUMENT ID
<b>0.499308<math>\pm</math>0.000005 OUR EVALUATION</b>	

## CP VIOLATION PARAMETERS in $B_s^0$

$$\text{Re}(\epsilon_{B_s^0}) / (1 + |\epsilon_{B_s^0}|^2)$$

CP impurity in  $B_s^0$  system.

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/scaling procedure takes into account correlation between the measurements. The value has been obtained from a 2D fit of the  $B_d$  and  $B_s$  asymmetries, which includes the  $B_s$  measurements listed below and the  $B$  factory average for the  $B_d$ .

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>−1.9 <math>\pm</math> 1.0 OUR EVALUATION</b>			
<b>−1.5 <math>\pm</math> 1.0 OUR AVERAGE</b>			
$-0.15 \pm 1.25 \pm 0.90$	1	AAIJ	14D LHCb $pp$ at 7 TeV
$-2.15 \pm 1.85$	2	ABAZOV	14 D0 $p\bar{p}$ at 1.96 TeV
$-2.8 \pm 1.9 \pm 0.4$	3	ABAZOV	13 D0 $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-4.5 \pm 2.7$	4	ABAZOV	11U D0 Repl. by ABAZOV 14
$-0.4 \pm 2.3 \pm 0.4$	5	ABAZOV	10E D0 Repl. by ABAZOV 13
$-3.6 \pm 1.9$	6	ABAZOV	10H D0 Repl. by ABAZOV 11U
$6.1 \pm 4.8 \pm 0.9$	7	ABAZOV	07A D0 Repl. by ABAZOV 10E

1 AAIJ 14D reports a measurement of time-integrated flavor-specific asymmetry in  $B_s^0 \rightarrow \mu^+ D_s^- X$  decays  $a_{SL}^s = (-0.06 \pm 0.50 \pm 0.36)\%$  which is approximately equal to  $4 \times \text{Re}(\epsilon_{B_s^0}) / (1 + |\epsilon_{B_s^0}|^2)$ .

2 ABAZOV 14 uses the dimuon charge asymmetry with different impact parameters from which it reports  $A_{SL}^s = (-0.86 \pm 0.74) \times 10^{-2}$ .

3 ABAZOV 13 reports a measurement of time-integrated flavor-specific asymmetry in mixed semileptonic  $B_s^0 \rightarrow \mu^+ D_s^- X$  decays  $A_{SL}^s = (-1.12 \pm 0.74 \pm 0.17)\%$  which is approximately equal to  $4 \times \text{Re}(\epsilon_{B_s^0}) / (1 + |\epsilon_{B_s^0}|^2)$ .

4 ABAZOV 11U uses the dimuon charge asymmetry with different impact parameters from which it reports  $A_{SL}^s = (-18.1 \pm 10.6) \times 10^{-3}$ .

5 ABAZOV 10E reports a measurement of flavor-specific asymmetry in  $B_{(s)}^0 \rightarrow \mu^+ D_{(s)}^{*-} X$  decays with a decay-time analysis including initial-state flavor tagging,  $A_{SL}^s = (-1.7 \pm 9.1^{+1.4}_{-1.5}) \times 10^{-3}$  which is approximately equal to  $4 \times \text{Re}(\epsilon_{B_s^0}) / (1 + |\epsilon_{B_s^0}|^2)$ .

# Meson Particle Listings

## $B_S^0$

<sup>6</sup>ABAZOV 10H reports a measurement of like-sign dimuon charge asymmetry of  $A_{SL}^b=(-9.57\pm 2.51\pm 1.46)\times 10^{-3}$  in semileptonic  $b$ -hadron decays. Using the measured production ratio of  $B_d^0$  and  $B_S^0$ , and the asymmetry of  $B_d^0$ ,  $A_{SL}^d=(-4.7\pm 4.6)\times 10^{-3}$  measured from  $B$ -factories, they obtain the asymmetry for  $B_S^0$ .

<sup>7</sup>The first direct measurement of the time integrated flavor untagged charge asymmetry in semileptonic  $B_S^0$  decays is reported as  $2\chi A_{SL}^s(\text{untagged})=A_{SL}^s=(2.45\pm 1.93\pm 0.35)\times 10^{-2}$ .

### $C_{KK}(B_S^0\rightarrow K^+K^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.14\pm 0.11\pm 0.03</math></b>	AAIJ	13B0 LHCB	$pp$ at 7 TeV

### $S_{KK}(B_S^0\rightarrow K^+K^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.30\pm 0.12\pm 0.04</math></b>	AAIJ	13B0 LHCB	$pp$ at 7 TeV

## $\gamma$

For angle  $\gamma(\phi_3)$  of the CKM unitarity triangle, see the review on “ $CP$  Violation” in the Reviews section.

VALUE(°)	DOCUMENT ID	TECN	COMMENT
<b><math>65\pm 7</math> OUR AVERAGE</b>			
$63.5^{+7.2}_{-6.7}$	<sup>1,2</sup> AAIJ	15K LHCB	$pp$ at 7, 8 TeV
$115^{+28}_{-43}$	<sup>3</sup> AAIJ	14Bf LHCB	$pp$ at 7 TeV

- <sup>1</sup> Obtained by measuring time-dependent  $CP$  asymmetry in  $B_S^0\rightarrow K^+K^-$  and using a U-spin relation between  $B_S^0\rightarrow K^+K^-$  and  $B^0\rightarrow \pi^+\pi^-$ .
- <sup>2</sup> Results are also presented using additional inputs on  $B^0\rightarrow \pi^0\pi^0$  and  $B^+\rightarrow \pi^+\pi^0$  decays from other experiments and isospin symmetry assumptions. The dependence of the results on the maximum allowed amount of U-spin breaking up to 50% is also included.
- <sup>3</sup> Measured in  $B_S^0\rightarrow D_S^\mp K^\pm$  decays, constraining  $-2\beta_S$  by the measurement of  $\phi_S=0.01\pm 0.07\pm 0.0$  from AAIJ 13AR. The value is modulo  $180^\circ$  at 68% CL.

### $\delta_B(B_S^0\rightarrow D_S^\pm K^\mp)$

VALUE(°)	DOCUMENT ID	TECN	COMMENT
<b><math>3^{+19}_{-20}</math></b>	<sup>1</sup> AAIJ	14Bf LHCB	$pp$ at 7 TeV

- <sup>1</sup> Measured in  $B_S^0\rightarrow D_S^\mp K^\pm$  decays, constraining  $-2\beta_S$  by the measurement of  $\phi_S=0.01\pm 0.07\pm 0.0$  from AAIJ 13AR. The value is modulo  $180^\circ$  at 68% CL.

### $r_B(B_S^0\rightarrow D_S^\mp K^\pm)$

$r_B$  and  $\delta_B$  are the amplitude ratio and relative strong phase between the amplitudes of  $A(B_S^0\rightarrow D_S^+K^-)$  and  $A(B_S^0\rightarrow D_S^-K^+)$ ,

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.53^{+0.17}_{-0.16}</math></b>	<sup>1</sup> AAIJ	14Bf LHCB	$pp$ at 7 TeV

- <sup>1</sup> Measured in  $B_S^0\rightarrow D_S^\mp K^\pm$  decays, constraining  $-2\beta_S$  by the measurement of  $\phi_S=0.01\pm 0.07\pm 0.0$  from AAIJ 13AR. At 68% CL.

## $CP$ Violation phase $\beta_s$

$-2\beta_s$  is the weak phase difference between  $B_S^0$  mixing amplitude and the  $B_S^0\rightarrow J/\psi\phi$  decay amplitude driven by the  $b\rightarrow c\overline{c}s$  transition (such as  $B_S\rightarrow J/\psi\phi$ ,  $J/\psi K^+K^-$ ,  $J/\psi\pi^+\pi^-$ , and  $D_S^+D_S^-$ ). The Standard Model value of  $\beta_s$  is  $\arg(-\frac{V_{ts}V_{cb}^*}{V_{cs}V_{tb}^*})$  if penguin contributions are neglected.

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/scaling procedure takes into account correlation between the measurements.

VALUE(10 <sup>-2</sup> rad)	DOCUMENT ID	TECN	COMMENT
<b><math>0.6\pm 1.9</math> OUR EVALUATION</b>			
<b><math>1.1\pm 1.9</math> OUR AVERAGE</b>			
$2.9\pm 2.5\pm 0.3$	<sup>1</sup> AAIJ	15i LHCB	$pp$ at 7, 8 TeV
$6^{+8}_{-7}$	<sup>2,3</sup> AAIJ	15K LHCB	$pp$ at 7, 8 TeV
$-6\pm 13\pm 3$	<sup>4</sup> AAD	14U ATLS	$pp$ at 7 TeV
$-1\pm 9\pm 1$	<sup>5</sup> AAIJ	14AY LHCB	$pp$ at 7, 8 TeV
$-3.5\pm 3.4\pm 0.4$	<sup>6</sup> AAIJ	14s LHCB	$pp$ at 7, 8 TeV
	<sup>7</sup> AALTONEN	12AJ CDF	$p\overline{p}$ at 1.96 TeV
$28^{+18}_{-19}$	<sup>8,9</sup> ABAZOV	12D D0	$p\overline{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-17\pm 15\pm 3$	<sup>10</sup> AAIJ	14AE LHCB	$pp$ at 7, 8 TeV
$-0.5\pm 3.5\pm 0.5$	<sup>11</sup> AAIJ	13AR LHCB	Repl. by AAIJ 15i
	<sup>12</sup> AAIJ	13AY LHCB	$pp$ at 7 TeV
$-11.0\pm 20.5\pm 5.0$	<sup>13</sup> AAD	12CV ATLS	Repl. by AAD 14U
$22\pm 22\pm 1$	<sup>14</sup> AAIJ	12B LHCB	Repl. by AAIJ 12Q
$-8\pm 9\pm 3$	<sup>15</sup> AAIJ	12D LHCB	Repl. by AAIJ 13AR

$0.95^{+8.70+0.15}_{-8.65-0.20}$	<sup>16</sup> AAIJ	12Q LHCB	Repl. by AAIJ 13AR
	<sup>17</sup> AALTONEN	12D CDF	Repl. by AALTONEN 12AJ
	<sup>18</sup> AALTONEN	08G CDF	Repl. by AALTONEN 12D
$28^{+12}_{-15}^{+4}_{-1}$	<sup>8,19</sup> ABAZOV	08AM D0	Repl. by ABAZOV 12D
$39.5\pm 28.0^{+0.5}_{-7.0}$	<sup>9,20</sup> ABAZOV	07 D0	Repl. by ABAZOV 07N
$35^{+20}_{-24}$	<sup>9,21</sup> ABAZOV	07N D0	Repl. by ABAZOV 08AM

- <sup>1</sup> AAIJ 15i reports  $\phi_s=-2\beta_s=-0.058\pm 0.049\pm 0.006$  rad. that was measured using a tagged, time-dependent angular analysis of  $B_S^0\rightarrow J/\psi K^+K^-$  decays. It also combines this result with that of AAIJ 14s and quotes  $\phi_s=-2\beta_s=-0.010\pm 0.039$  rad.
- <sup>2</sup> AAIJ 15K reports  $-2\beta_s=-0.12^{+0.14}_{-0.16}$  rad. The value was obtained by measuring time-dependent  $CP$  asymmetry in  $B_S^0\rightarrow K^+K^-$  and using a U-spin relation between  $B_S^0\rightarrow K^+K^-$  and  $B^0\rightarrow \pi^+\pi^-$ .
- <sup>3</sup> Results are also presented using additional inputs on  $B^0\rightarrow \pi^0\pi^0$  and  $B^+\rightarrow \pi^+\pi^0$  decays from other experiments and isospin symmetry assumptions. The dependence of the results on the maximum allowed amount of U-spin breaking up to 50% is also included.
- <sup>4</sup> AAD 14U reports  $\phi_s=-2\beta_s=0.12\pm 0.25\pm 0.05$  rad. that was measured using a tagged, time-dependent angular analysis of  $B_S^0\rightarrow J/\psi\phi$  decays.
- <sup>5</sup> AAIJ 14AY reports  $-2\beta_s=0.02\pm 0.17\pm 0.02$  rad. in tagged, time-dependent fit to  $B_S^0\rightarrow D_S^+D_S^-$ , while allowing  $CP$  violation in decay.
- <sup>6</sup> AAIJ 14s reports  $\phi_s=-2\beta_s=0.070\pm 0.068\pm 0.008$  rad. and  $|\lambda|=0.89\pm 0.05\pm 0.01$ , when direct  $CP$  violation is allowed. Measured using a tagged, time-dependent fit to  $B_S^0\rightarrow J/\psi\pi^+\pi^-$  decays.
- <sup>7</sup> AALTONEN 12AJ reports  $-\pi/2<\beta_s<-1.51$  or  $-0.06<\beta_s<0.30$ , or  $1.26<\beta_s<\pi/2$  rad. at 68% CL. Measured using the time-dependent angular analysis of  $B_S^0\rightarrow J/\psi\phi$  decays.
- <sup>8</sup> Measured using fully reconstructed  $B_S\rightarrow J/\psi\phi$  decays. A single error includes both statistical and systematic uncertainties.
- <sup>9</sup> Reports  $\phi_s$  which equals to  $-2\beta_s$ .
- <sup>10</sup> Measured in  $B_S^0\rightarrow \phi\phi$  decays. This is a  $b\rightarrow s\overline{s}s$  transition with a decay amplitude phase different from that of  $b\rightarrow c\overline{c}s$  transition.
- <sup>11</sup> AAIJ 13AR reports  $\phi_s=-2\beta_s=0.01\pm 0.07\pm 0.01$  rad. obtained from combined fit to  $B_S^0\rightarrow J/\psi K^+K^-$  and  $B_S^0\rightarrow J/\psi\pi^+\pi^-$  data sets. Also reports separate results of  $\phi_s=0.07\pm 0.09\pm 0.01$  rad. from  $B_S^0\rightarrow J/\psi K^+K^-$  decays and  $\phi_s=-0.14^{+0.17}_{-0.16}\pm 0.01$  rad. from  $B_S^0\rightarrow J/\psi\pi^+\pi^-$  decays.
- <sup>12</sup> AAIJ 13AY uses  $B_S^0\rightarrow \phi\phi$  mode, and reports the 68% CL interval of  $\phi_s=-2\beta_s$  as  $[-2.46,-0.76]$  rad.
- <sup>13</sup> AAD 12CV reports  $\phi_s=-2\beta_s=0.22\pm 0.41\pm 0.10$  rad. that was measured using a time-dependent angular analysis of  $B_S^0\rightarrow J/\psi\phi$  decays.
- <sup>14</sup> Reports  $\phi_s=-2\beta_s=-0.44\pm 0.44\pm 0.02$  rad. that was measured using a time-dependent fit to  $B_S^0\rightarrow J/\psi f_0(980)$  decays.
- <sup>15</sup> Reports  $\phi_s=-2\beta_s=0.15\pm 0.18\pm 0.06$  rad. that was measured using a time-dependent angular analysis of  $B_S^0\rightarrow J/\psi\phi$  decays.
- <sup>16</sup> Reports  $\phi_s=-2\beta_s=-0.019^{+0.173+0.004}_{-0.174-0.003}$  rad. which was measured using a time-dependent fit to  $B_S^0\rightarrow J/\psi\pi^+\pi^-$  decays, with the  $\pi^+\pi^-$  mass within 775–1550 MeV. Searches for, but finds no evidence, for direct  $CP$  violation in  $B_S^0\rightarrow J/\psi\pi\pi$  decays.
- <sup>17</sup> Reports  $0.02<\phi_s<0.52$  or  $1.08<\phi_s<1.55$  rad. at 68% C.L. confidence regions in the two-dimensional space of  $\phi_s$  and  $\Delta\Gamma_{B_S^0}$  from  $B_S^0\rightarrow J/\psi\phi$  decays.
- <sup>18</sup> Reports  $0.32<2\beta_s<2.82$  rad. at 68% C.L. and confidence regions in the two-dimensional space of  $2\beta_s$  and  $\Delta\Gamma$  from the first measurement of  $B^0\rightarrow J/\psi\phi$  decays using flavor tagging. The probability of a deviation from SM prediction as large as the level of observed data is 15%.
- <sup>19</sup> Reports  $\phi_s=-2\beta_s$  and obtains 90% CL interval  $-0.03<\beta_s<0.60$  rad.
- <sup>20</sup> The first direct measurement of the  $CP$ -violating mixing phase is reported from the time-dependent analysis of flavor untagged  $B_S^0\rightarrow J/\psi\phi$  decays.
- <sup>21</sup> Combines D0 collaboration measurements of time-dependent angular distributions in  $B_S^0\rightarrow J/\psi\phi$  and charge asymmetry in semileptonic decays. There is a 4-fold ambiguity in the solution.

### $|\lambda|(B_S^0\rightarrow J/\psi(1S)\phi)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.964\pm 0.019\pm 0.007</math></b>	AAIJ	15i LHCB	$pp$ at 7, 8 TeV

## $|\lambda|$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.02\pm 0.07</math> OUR AVERAGE</b>			
$1.04\pm 0.07\pm 0.03$	<sup>1</sup> AAIJ	14AE LHCB	$pp$ at 7, 8 TeV
$0.91^{+0.18}_{-0.15}\pm 0.02$	<sup>2</sup> AAIJ	14AY LHCB	$pp$ at 7, 8 TeV

- <sup>1</sup> Measured in  $B_S^0\rightarrow \phi\phi$  decays.
- <sup>2</sup> Measured in  $B_S^0\rightarrow D_S^+D_S^-$  decays.

## A, $CP$ violation parameter

$A=-2\text{Re}(\lambda)/(1+|\lambda|^2)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.49^{+0.77}_{-0.65}\pm 0.06</math></b>	<sup>1</sup> AAIJ	15AL LHCB	$pp$ at 7, 8 TeV

- <sup>1</sup> Measured in  $B_S^0\rightarrow J/\psi K_S^0$  decays.

See key on page 601

## Meson Particle Listings

 $B_s^0$ 

## C, CP violation parameter

$$C = (1 - |\lambda|^2) / (1 + |\lambda|^2)$$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.28 \pm 0.41 \pm 0.08$	<sup>1</sup> AAIJ	15AL LHCb	$p\bar{p}$ at 7, 8 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow J/\psi K_S^0$ decays.			

## S, CP violation parameter

$$S = -2 \operatorname{Im}(\lambda) / (1 + |\lambda|^2)$$

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08 \pm 0.40 \pm 0.08$	<sup>1</sup> AAIJ	15AL LHCb	$p\bar{p}$ at 7, 8 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow J/\psi K_S^0$ decays.			

 $A_{CP}^L(B_s \rightarrow J/\psi \bar{K}^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.048 \pm 0.057 \pm 0.020$	AAIJ	15AV LHCb	$p\bar{p}$ at 7, 8 TeV

 $A_{CP}^{\parallel}(B_s \rightarrow J/\psi \bar{K}^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$0.171 \pm 0.152 \pm 0.028$	AAIJ	15AV LHCb	$p\bar{p}$ at 7, 8 TeV

 $A_{CP}^{\perp}(B_s \rightarrow J/\psi \bar{K}^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.049 \pm 0.096 \pm 0.025$	AAIJ	15AV LHCb	$p\bar{p}$ at 7, 8 TeV

 $A_{CP}(B_s \rightarrow \pi^+ K^-)$  $A_{CP}$  is defined as

$$\frac{B(\bar{B}^0 \rightarrow f) - B(B^0 \rightarrow \bar{f})}{B(\bar{B}_s^0 \rightarrow f) + B(B_s^0 \rightarrow \bar{f})},$$

the CP-violation asymmetry of exclusive  $B_s^0$  and  $\bar{B}_s^0$  decay.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.263 \pm 0.035</math> OUR AVERAGE</b>			
$0.22 \pm 0.07 \pm 0.02$	AALTONEN	14P CDF	$p\bar{p}$ at 1.96 TeV
$0.27 \pm 0.04 \pm 0.01$	AAIJ	13AX LHCb	$p\bar{p}$ at 7 TeV
$0.39 \pm 0.15 \pm 0.08$	AALTONEN	11N CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.27 \pm 0.08 \pm 0.02$	AAIJ	12V LHCb	Repl. by AAIJ 13AX

 $A_{CP}(B_s^0 \rightarrow [K^+ K^-]_D \bar{K}^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.04 \pm 0.07 \pm 0.02$	AAIJ	14BN LHCb	$p\bar{p}$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.04 \pm 0.16 \pm 0.01$	AAIJ	13L LHCb	Repl. by AAIJ 14BN

 $A_{CP}(B_s^0 \rightarrow [\pi^+ K^-]_D K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.01 \pm 0.03 \pm 0.02$	AAIJ	14BN LHCb	$p\bar{p}$ at 7, 8 TeV

 $A_{CP}(B_s^0 \rightarrow [\pi^+ \pi^-]_D K^*(892)^0)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.13 \pm 0.02</math></b>	AAIJ	14BN LHCb	$p\bar{p}$ at 7, 8 TeV

## CPT VIOLATION PARAMETERS

In the  $B_s^0$  mixing, propagating mass eigenstates can be written as

$$\begin{aligned} |B_{sL}\rangle &\propto p \sqrt{1-\xi} |B_s^0\rangle + q \sqrt{1+\xi} |\bar{B}_s^0\rangle \\ |B_{sH}\rangle &\propto p \sqrt{1+\xi} |B_s^0\rangle - q \sqrt{1-\xi} |\bar{B}_s^0\rangle \end{aligned}$$

where parameter  $\xi$  controls  $CPT$  violation. If  $\xi$  is zero, then  $CPT$  is conserved. The parameter  $\xi$  can be written as

$$\xi = \frac{2(M_{11}-M_{22})-i(\Gamma_{11}-\Gamma_{22})}{-2\Delta m_s + i\Delta\Gamma_s} \approx \frac{-2\beta^\mu \Delta a_\mu}{2\Delta m_s - i\Delta\Gamma_s},$$

where  $M_{ii}$ ,  $\Gamma_{ii}$ ,  $\Delta m_s$ , and  $\Delta\Gamma_s$  are parameters of Hamiltonian governing  $B_s^0$  oscillations,  $\beta^\mu$  is the  $B_s^0$  meson velocity and  $\Delta a_\mu$  characterizes Lorentz-invariance violation. $\Delta a_\perp$ 

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 1.2 \times 10^{-12}</math></b>	95	<sup>1</sup> ABAZOV	15L D0	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Measured in semileptonic $B_s^0 \rightarrow D_s^- \mu^+ X$ decays. Also extracts limit on time and longitudinal components ( $-0.8 < \Delta a_T - 0.396 \Delta a_Z < 3.9$ ) $10^{-13}$ GeV.				

PARTIAL BRANCHING FRACTIONS IN  $B_s \rightarrow \phi \ell^+ \ell^-$  $B(B_s \rightarrow \phi \ell^+ \ell^-) (0.1 < q^2 < 2.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.14 \pm 0.16</math> OUR AVERAGE</b>			
$1.11^{+0.14}_{-0.13} \pm 0.09$	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
$2.78 \pm 0.95 \pm 0.89$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.897^{+0.207}_{-0.186} \pm 0.097$	<sup>1</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (2.0 < q^2 < 5.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.77 \pm 0.12 \pm 0.06</math></b>	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.529^{+0.182}_{-0.159} \pm 0.057$	<sup>1,2</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
$0.58 \pm 0.55 \pm 0.19$	<sup>2</sup> AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			
<sup>2</sup> Measured in $2 < q^2 < 4.3 \text{ GeV}^2/c^4$ .			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (5.0 < q^2 < 8.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.96 \pm 0.13 \pm 0.08</math></b>	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.38^{+0.25}_{-0.23} \pm 0.14$	<sup>1,2</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
$1.34 \pm 0.83 \pm 0.43$	<sup>2</sup> AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			
<sup>2</sup> Measured in $4.3 < q^2 < 8.68 \text{ GeV}^2/c^4$ .			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (11.0 < q^2 < 12.5 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.71 \pm 0.10 \pm 0.06</math></b>	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.18^{+0.22}_{-0.21} \pm 0.14$	<sup>1,2</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
$2.98 \pm 0.95 \pm 0.95$	<sup>2</sup> AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			
<sup>2</sup> Measured in $10.9 < q^2 < 12.86 \text{ GeV}^2/c^4$ .			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (15.0 < q^2 < 17.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.90 \pm 0.11 \pm 0.07</math></b>	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.760^{+0.189}_{-0.169} \pm 0.087$	<sup>1,2</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
$1.86 \pm 0.66 \pm 0.59$	<sup>2</sup> AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			
<sup>2</sup> Measured in $14.18 < q^2 < 16 \text{ GeV}^2/c^4$ .			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (17.0 < q^2 < 19.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.79 \pm 0.11 \pm 0.07</math></b>	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.06^{+0.23}_{-0.21} \pm 0.12$	<sup>1,2</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
$2.32 \pm 0.76 \pm 0.74$	<sup>2</sup> AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			
<sup>2</sup> Measured in $16 < q^2 < 19 \text{ GeV}^2/c^4$ .			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.28 \pm 0.18</math> OUR AVERAGE</b>			
$1.29 \pm 0.16 \pm 0.10$	<sup>1</sup> AAIJ	15AQ LHCb	$p\bar{p}$ at 7, 8 TeV
$1.14 \pm 0.79 \pm 0.36$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.14^{+0.25}_{-0.23} \pm 0.13$	<sup>1</sup> AAIJ	13X LHCb	Repl. by AAIJ 15AQ
<sup>1</sup> Measured in $B_s^0 \rightarrow \phi \mu^+ \mu^-$ decays.			

 $B(B_s \rightarrow \phi \ell^+ \ell^-) (0.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$ 

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.30 \pm 1.09 \pm 1.05</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV



See key on page 601

Meson Particle Listings

$B_s^0, B_s^*, B_{s1}(5830)^0, B_{s2}^*(5840)^0$

AKERS	95G	PL B350 273	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95J	ZPHY C66 555	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	95J	PL B356 409	D. Buskulić <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	95O	PL B361 221	D. Buskulić <i>et al.</i>	(ALEPH Collab.)
ABREU	94D	PL B324 500	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	94E	ZPHY C61 407	P. Abreu <i>et al.</i>	(DELPHI Collab.)
Also		PL B289 199	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	94J	PL B337 196	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	94L	PL B337 393	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	94B	PL B322 441	D. Buskulić <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94C	PL B322 275	D. Buskulić <i>et al.</i>	(ALEPH Collab.)
ABE	93F	PRL 71 1685	F. Abe <i>et al.</i>	(CDF Collab.)
ACTON	93H	PL B312 501	P.D. Acton <i>et al.</i>	(OPAL Collab.)
BUSKULIC	93G	PL B311 425	D. Buskulić <i>et al.</i>	(ALEPH Collab.)
ABREU	92M	PL B289 199	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	92N	PL B295 357	P.D. Acton <i>et al.</i>	(OPAL Collab.)
BUSKULIC	92E	PL B294 145	D. Buskulić <i>et al.</i>	(ALEPH Collab.)
LEE-FRANZINI	90	PRL 65 2947	J. Lee-Franzini <i>et al.</i>	(CUSP II Collab.)

$B_s^*$

$I(J^P) = 0(1^-)$

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

$B_s^*$  MASS

From mass difference below and the  $B_s^0$  mass.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5415.4<math>^{+1.8}_{-1.5}</math> OUR FIT</b>			Error includes scale factor of 3.0.
<b>5415.8<math>\pm 1.5</math> OUR AVERAGE</b>			Error includes scale factor of 2.6.
5416.4 $\pm 0.4 \pm 0.5$	LOUVOT 09	BELL	$e^+e^- \rightarrow \Upsilon(5S)$
5411.7 $\pm 1.6 \pm 0.6$	<sup>1</sup> AQUINES 06	CLEO	$e^+e^- \rightarrow \Upsilon(5S)$
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
5418 $\pm 1 \pm 3$	DRUTSKOY 07A	BELL	Repl. by LOUVOT 09
5414 $\pm 1 \pm 3$	<sup>2</sup> BONVICINI 06	CLEO	$e^+e^- \rightarrow \Upsilon(5S)$
<sup>1</sup> Utilized the beam constrained invariant mass peak positions for $B^*$ and $B_s^*$ to extract the measurement.			
<sup>2</sup> Uses 14 candidates consistent with $B_s$ decays into final states with a $J/\psi$ and a $D_s^{(*)-}$ .			

$m_{B_s^*} - m_{B_s}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>48.6<math>^{+1.8}_{-1.6}</math> OUR FIT</b>			Error includes scale factor of 2.8.
<b>46.1<math>\pm 1.5</math> OUR AVERAGE</b>			
45.7 $\pm 1.7 \pm 0.7$	<sup>3</sup> AQUINES 06	CLEO	$e^+e^- \rightarrow \Upsilon(5S)$
47.0 $\pm 2.6$	<sup>4</sup> LEE-FRANZINI 90	CSB2	$e^+e^- \rightarrow \Upsilon(5S)$
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
48 $\pm 1 \pm 3$	<sup>5</sup> BONVICINI 06	CLEO	Repl. by AQUINES 06
<sup>3</sup> Utilized the beam constrained invariant mass peak positions for $B^*$ and $B_s^*$ to extract the measurement.			
<sup>4</sup> LEE-FRANZINI 90 measure 46.7 $\pm 0.4 \pm 0.2$ MeV for an admixture of $B^0, B^+,$ and $B_s$ . They use the shape of the photon line to separate the above value for $B_s$ .			
<sup>5</sup> Uses 14 candidates consistent with $B_s$ decays into final states with a $J/\psi$ and a $D_s^{(*)-}$ .			

$|(m_{B_s^*} - m_{B_s}) - (m_{B^*} - m_B)|$

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6</b>	95	ABREU 95R	DLPH	$E_{\text{cm}}^{ee} = 88\text{--}94$ GeV

$B_s^*$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad B_s \gamma$	dominant

$B_s^*$  REFERENCES

LOUVOT 09	PRL 102 021801	R. Louvot <i>et al.</i>	(BELLE Collab.)
DRUTSKOY 07A	PR D76 012002	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
AQUINES 06	PRL 96 152001	O. Aquines <i>et al.</i>	(CLEO Collab.)
BONVICINI 06	PRL 96 022002	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
ABREU 95R	ZPHY C68 353	P. Abreu <i>et al.</i>	(DELPHI Collab.)
LEE-FRANZINI 90	PRL 65 2947	J. Lee-Franzini <i>et al.</i>	(CUSP II Collab.)

$B_{s1}(5830)^0$

$I(J^P) = 0(1^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

$B_{s1}(5830)^0$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5828.63<math>\pm 0.27</math> OUR FIT</b>			
<b>5828.40<math>\pm 0.04 \pm 0.41</math></b>	<sup>1</sup> AAIJ 13o	LHCB	$p\bar{p}$ at 7 TeV
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
5829.4 $\pm 0.7$	<sup>2</sup> AALTONEN 08k	CDF	Repl. by AALTONEN 14i
<sup>1</sup> Uses $B_{s1}(5830)^0 \rightarrow B^{*+} K^-$ decay.			
<sup>2</sup> Uses two-body decays into $K^-$ and $B^+$ mesons reconstructed as $B^+ \rightarrow J/\psi K^+, J/\psi \rightarrow \mu^+\mu^-$ or $B^+ \rightarrow \bar{D}^0 \pi^+, \bar{D}^0 \rightarrow K^+ \pi^-$ .			

$m_{B_{s1}^0} - m_{B^{*+}}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>503.98<math>\pm 0.18</math> OUR FIT</b>			
<b>504.03<math>\pm 0.12 \pm 0.15</math></b>	<sup>3</sup> AALTONEN 14i	CDF	$p\bar{p}$ at 1.96 TeV
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
504.41 $\pm 0.21 \pm 0.14$	<sup>4</sup> AALTONEN 08k	CDF	Repl. by AALTONEN 14i
<sup>3</sup> AALTONEN 14i reports $m_{B_{s1}(5830)^0} - m_{B^{*+}} - m_{K^-} = 10.35 \pm 0.12 \pm 0.15$ MeV which we adjusted by the $K^-$ mass.			
<sup>4</sup> Uses two-body decays into $K^-$ and $B^+$ mesons reconstructed as $B^+ \rightarrow J/\psi K^+, J/\psi \rightarrow \mu^+\mu^-$ or $B^+ \rightarrow \bar{D}^0 \pi^+, \bar{D}^0 \rightarrow K^+ \pi^-$ .			

$B_{s1}(5830)^0$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>0.5<math>\pm 0.3 \pm 0.3</math></b>	AALTONEN 14i	CDF	$p\bar{p}$ at 1.96 TeV

$B_{s1}(5830)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad B^{*+} K^-$	dominant

$B_{s1}(5830)^0$  BRANCHING RATIOS

$\Gamma(B^{*+} K^-)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
dominant	AALTONEN 08k	CDF	$p\bar{p}$ at 1.96 TeV

$B_{s1}(5830)^0$  REFERENCES

AALTONEN 14i	PR D90 012013	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AAIJ 13O	PRL 110 151803	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN 08K	PRL 100 082001	T. Aaltonen <i>et al.</i>	(CDF Collab.)

$B_{s2}^*(5840)^0$

$I(J^P) = 0(2^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

$B_{s2}^*(5840)^0$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5839.84<math>\pm 0.18</math> OUR FIT</b>			Error includes scale factor of 1.1.
<b>5839.98<math>\pm 0.20</math> OUR AVERAGE</b>			
5839.99 $\pm 0.05 \pm 0.20$	AAIJ 13o	LHCB	$p\bar{p}$ at 7 TeV
5839.6 $\pm 1.1 \pm 0.7$	<sup>1</sup> ABAZOV 08E	D0	$p\bar{p}$ at 1.96 TeV
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
5839.7 $\pm 0.7$	<sup>2</sup> AALTONEN 08k	CDF	Repl. by AALTONEN 14i
<sup>1</sup> Observed in $B_{s2}^{*0} \rightarrow B^+ K^-$ . Measured production rate of $B_{s2}^{*0}$ relative to $B^+$ to be (1.15 $\pm 0.23 \pm 0.13$ )%.			
<sup>2</sup> Uses two-body decays into $K^-$ and $B^+$ mesons reconstructed as $B^+ \rightarrow J/\psi K^+, J/\psi \rightarrow \mu^+\mu^-$ or $B^+ \rightarrow \bar{D}^0 \pi^+, \bar{D}^0 \rightarrow K^+ \pi^-$ .			

$m_{B_{s2}^0} - m_{B_{s1}^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
10.5 $\pm 0.6$	<sup>3</sup> AALTONEN 08k	CDF	Repl. by AALTONEN 14i
<sup>3</sup> Uses two-body decays into $K^-$ and $B^+$ mesons reconstructed as $B^+ \rightarrow J/\psi K^+, J/\psi \rightarrow \mu^+\mu^-$ or $B^+ \rightarrow \bar{D}^0 \pi^+, \bar{D}^0 \rightarrow K^+ \pi^-$ .			

Meson Particle Listings

$B_{s2}^*(5840)^0, B_{sJ}^*(5850)$

$m_{B_{s2}^0} - m_{B^+}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>560.53±0.18 OUR FIT</b>	Error includes scale factor of 1.1.		
<b>560.41±0.13±0.14</b>	<sup>4</sup> AALTONEN	14i CDF	$p\overline{p}$ at 1.96 TeV
<sup>4</sup> AALTONEN 14i reports $m_{B_{s2}(5840)^0} - m_{B^+} - m_{K^-} = 66.73 \pm 0.13 \pm 0.14$ MeV which we adjusted by the $K^-$ mass.			

$B_{s2}^*(5840)^0$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1.47±0.33 OUR AVERAGE</b>			
1.4 ±0.4 ±0.2	AALTONEN	14i CDF	$p\overline{p}$ at 1.96 TeV
1.56±0.13±0.47	<sup>5</sup> AAIJ	13o LHCB	$p\overline{p}$ at 7 TeV
<sup>5</sup> Uses $B_{s2}^*(5840)^0 \rightarrow B^{*+} K^-$ decays.			

$B_{s2}^*(5840)^0$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $B^+ K^-$	dominant
$\Gamma_2$ $B^{*+} K^-$	

$B_{s2}^*(5840)^0$ BRANCHING RATIOS			
$\Gamma(B^+ K^-)/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT
<b>dominant</b>	AALTONEN	08k CDF	$p\overline{p}$ at 1.96 TeV
<b>dominant</b>	<sup>6</sup> ABAZOV	08E D0	$p\overline{p}$ at 1.96 TeV
<sup>6</sup> Measured production rate of $B_{s2}^{*0}$ relative to $B^+$ to be $(1.15 \pm 0.23 \pm 0.13)\%$ .			

$\Gamma(B^{*+} K^-)/\Gamma(B^+ K^-)$			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.093±0.013±0.012</b>	AAIJ	13o LHCB	$p\overline{p}$ at 7 TeV

$B_{s2}^*(5840)^0$ REFERENCES			
AALTONEN	14i	PR D90 012013	T. Aaltonen <i>et al.</i> (CDF Collab.)
AAIJ	13o	PRL 110 151803	R. Aaij <i>et al.</i> (LHCb Collab.)
AALTONEN	08k	PRL 100 082001	T. Aaltonen <i>et al.</i> (CDF Collab.)
ABAZOV	08E	PRL 100 082002	V.M. Abazov <i>et al.</i> (D0 Collab.)

$B_{sJ}^*(5850)$

$I(J^P) = ?(?^?)$   
 $I, J, P$  need confirmation.

OMITTED FROM SUMMARY TABLE

Signal can be interpreted as coming from  $\overline{b}s$  states. Needs confirmation.

$B_{sJ}^*(5850)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5853±15</b>	14i	AKERS	95E OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

$B_{sJ}^*(5850)$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>47±22</b>	14i	AKERS	95E OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94$ GeV

$B_{sJ}^*(5850)$ REFERENCES				
AKERS	95E	ZPHY C66 19	R. Akers <i>et al.</i>	(OPAL Collab.)



See key on page 601

## Meson Particle Listings

 $B_c^+$ 

# BOTTOM, CHARMED MESONS ( $B = C = \pm 1$ )

 $B_c^+ = c\bar{b}$ ,  $B_c^- = \bar{c}b$ , similarly for  $B_c^{* \pm}$ 's $B_c^+$ 
 $I(J^P) = 0(0^-)$   
 $I, J, P$  need confirmation.

Quantum numbers shown are quark-model predictions.

 $B_c^+$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>6275.1 <math>\pm</math> 1.0 OUR AVERAGE</b>			
6274.0 $\pm$ 1.8 $\pm$ 0.4	<sup>1</sup> AAIJ	14AQ LHCb	$p\bar{p}$ at 7, 8 TeV
6276.28 $\pm$ 1.44 $\pm$ 0.36	<sup>2</sup> AAIJ	13AS LHCb	$p\bar{p}$ at 7, 8 TeV
6273.7 $\pm$ 1.3 $\pm$ 1.6	<sup>3</sup> AAIJ	12AV LHCb	$p\bar{p}$ at 7 TeV
6275.6 $\pm$ 2.9 $\pm$ 2.5	<sup>4</sup> AALTONEN	08M CDF	$p\bar{p}$ at 1.96 TeV
6300 $\pm$ 14 $\pm$ 5	<sup>4</sup> ABAZOV	08T D0	$p\bar{p}$ at 1.96 TeV
6400 $\pm$ 390 $\pm$ 130	<sup>5</sup> ABE	98M CDF	$p\bar{p}$ at 1.8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6285.7 $\pm$ 5.3 $\pm$ 1.2	<sup>4</sup> ABULENCIA	06C CDF	Repl. by AALTONEN 08M
6320 $\pm$ 60	<sup>6</sup> ACKERSTAFF	98O OPAL	$e^+e^- \rightarrow Z$

<sup>1</sup> Uses  $B_c^+ \rightarrow J/\psi p \bar{p} \pi^+$  decays.<sup>2</sup> AAIJ 13AS uses the  $B_c^+ \rightarrow J/\psi D_s^+$ .<sup>3</sup> AAIJ 12AV uses the  $B_c^+ \rightarrow J/\psi \pi^+$  mode and also measures the mass difference  $M(B_c^+) - M(B^+) = 994.6 \pm 1.3 \pm 0.6$  MeV/ $c^2$ .<sup>4</sup> Measured using a fully reconstructed decay mode of  $B_c \rightarrow J/\psi \pi$ .<sup>5</sup> ABE 98M observed  $20.4^{+6.2}_{-5.5}$  events in the  $B_c^+ \rightarrow J/\psi(1S) \ell \nu_\ell$  with a significance of  $> 4.8$  standard deviations. The mass value is estimated from  $m(J/\psi(1S) \ell)$ .<sup>6</sup> ACKERSTAFF 98O observed 2 candidate events in the  $B_c^+ \rightarrow J/\psi(1S) \pi^+$  channel with an estimated background of  $0.63 \pm 0.20$  events. $B_c^+$  MEAN LIFE

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements.

VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
<b>0.507 <math>\pm</math> 0.009 OUR EVALUATION</b>			
<b>0.507 <math>\pm</math> 0.009 OUR AVERAGE</b>			
0.5134 $\pm$ 0.0110 $\pm$ 0.0057	<sup>1,2</sup> AAIJ	15G LHCb	$p\bar{p}$ at 7, 8 TeV
0.509 $\pm$ 0.008 $\pm$ 0.012	<sup>3</sup> AAIJ	14G LHCb	$p\bar{p}$ at 8 TeV
0.452 $\pm$ 0.048 $\pm$ 0.027	<sup>2</sup> AALTONEN	13 CDF	$p\bar{p}$ at 1.96 TeV
0.448 $\pm$ 0.038 $\pm$ 0.032	<sup>4</sup> ABAZOV	09H D0	$p\bar{p}$ at 1.96 TeV
0.463 $\pm$ 0.073 $\pm$ 0.036	<sup>4</sup> ABULENCIA	06O CDF	$p\bar{p}$ at 1.96 TeV
0.46 $\pm$ 0.18 $\pm$ 0.03	<sup>4</sup> ABE	98M CDF	$p\bar{p}$ 1.8 TeV

<sup>1</sup> Also measures the width difference  $\Delta\Gamma = \Gamma_{B_c^+} - \Gamma_{B^+} = 4.46 \pm 0.14 \pm 0.07$  mm $^{-1}$  c.<sup>2</sup> Uses fully reconstructed  $B_c^+ \rightarrow J/\psi \pi^+$  decays.<sup>3</sup> Measured using  $B_c^+ \rightarrow J/\psi \mu^+ \nu_\mu X$  decays.<sup>4</sup> The lifetime is measured from the  $J/\psi e$  decay vertices. $B_c^+$  DECAY MODES  $\times B(\bar{B} \rightarrow B_c)$  $B_c^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
The following quantities are not pure branching ratios; rather the fraction $\Gamma_i/\Gamma \times B(\bar{B} \rightarrow B_c)$ .		
$\Gamma_1$ $J/\psi(1S) \ell^+ \nu_\ell \text{ anything}$	$(5.2^{+2.4}_{-2.1}) \times 10^{-5}$	
$\Gamma_2$ $J/\psi(1S) \mu^+ \nu_\mu$		
$\Gamma_3$ $J/\psi(1S) \pi^+$	seen	
$\Gamma_4$ $J/\psi(1S) K^+$	seen	
$\Gamma_5$ $J/\psi(1S) \pi^+ \pi^+ \pi^-$	seen	
$\Gamma_6$ $J/\psi(1S) a_1(1260)$	$< 1.2$	$\times 10^{-3}$ 90%
$\Gamma_7$ $J/\psi(1S) K^+ K^- \pi^+$	seen	

$\Gamma_8$ $J/\psi(1S) \pi^+ \pi^+ \pi^- \pi^-$	seen	
$\Gamma_9$ $\psi(2S) \pi^+$	seen	
$\Gamma_{10}$ $J/\psi(1S) D_s^{*+}$	seen	
$\Gamma_{11}$ $J/\psi(1S) D_s^{*+}$	seen	
$\Gamma_{12}$ $J/\psi(1S) p \bar{p} \pi^+$	seen	
$\Gamma_{13}$ $D^*(2010)^+ \bar{D}^0$	$< 6.2$	$\times 10^{-3}$ 90%
$\Gamma_{14}$ $D^+ K^{*0}$	$< 0.20$	$\times 10^{-6}$ 90%
$\Gamma_{15}$ $D^+ \bar{K}^{*0}$	$< 0.16$	$\times 10^{-6}$ 90%
$\Gamma_{16}$ $D_s^+ K^{*0}$	$< 0.28$	$\times 10^{-6}$ 90%
$\Gamma_{17}$ $D_s^+ \bar{K}^{*0}$	$< 0.4$	$\times 10^{-6}$ 90%
$\Gamma_{18}$ $D_s^+ \phi$	$< 0.32$	$\times 10^{-6}$ 90%
$\Gamma_{19}$ $K^+ K^0$	$< 4.6$	$\times 10^{-7}$ 90%
$\Gamma_{20}$ $B_s^0 \pi^+ / B(\bar{B} \rightarrow B_s)$	$(2.37^{+0.37}_{-0.35}) \times 10^{-3}$	

 $B_c^+$  BRANCHING RATIOS $\Gamma(J/\psi(1S) \ell^+ \nu_\ell \text{ anything})/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$  $\Gamma_i/\Gamma \times B$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>(5.2^{+2.4}_{-2.1}) \times 10^{-5}</math></b>		<sup>1</sup> ABE	98M CDF	$p\bar{p}$ 1.8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.6$	$\times 10^{-4}$ 90	<sup>2</sup> ACKERSTAFF	98O OPAL	$e^+e^- \rightarrow Z$
$< 1.9$	$\times 10^{-4}$ 90	<sup>3</sup> ABREU	97E DLPH	$e^+e^- \rightarrow Z$
$< 1.2$	$\times 10^{-4}$ 90	<sup>4</sup> BARATE	97H ALEP	$e^+e^- \rightarrow Z$

<sup>1</sup> ABE 98M result is derived from the measurement of  $[\sigma(B_c) \times B(B_c \rightarrow J/\psi(1S) \ell \nu_\ell)] / [\sigma(B^+) \times B(B^+ \rightarrow J/\psi(1S) K^+)] = 0.132^{+0.041}_{-0.037}(\text{stat}) \pm 0.031(\text{sys}) \pm 0.032(\text{lifetime})$  by using PDG 98 values of  $B(b \rightarrow B^+)$  and  $B(B^+ \rightarrow J/\psi(1S) K^+)$ .<sup>2</sup> ACKERSTAFF 98O reports  $B(Z \rightarrow B_c X)/B(Z \rightarrow q\bar{q}) \times B(B_c \rightarrow J/\psi(1S) \ell \nu_\ell) < 6.95 \times 10^{-5}$  at 90%CL. We rescale to our PDG 98 values of  $B(Z \rightarrow b\bar{b})$ .<sup>3</sup> ABREU 97E value listed is for an assumed  $\tau_{B_c} = 0.4$  ps and improves to  $1.6 \times 10^{-4}$  for  $\tau_{B_c} = 1.4$  ps.<sup>4</sup> BARATE 97H reports  $B(Z \rightarrow B_c X)/B(Z \rightarrow q\bar{q}) \times B(B_c \rightarrow J/\psi(1S) \ell \nu_\ell) < 5.2 \times 10^{-5}$  at 90%CL. We rescale to our PDG 96 values of  $B(Z \rightarrow b\bar{b})$ . A  $B_c^+ \rightarrow J/\psi(1S) \mu^+ \nu_\mu$  candidate event is found, compared to all the known background sources  $2 \times 10^{-3}$ , which gives  $m_{B_c} = 5.96^{+0.25}_{-0.19}$  GeV and  $\tau_{B_c} = 1.77 \pm 0.17$  ps. $\Gamma(J/\psi(1S) \pi^+)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$  $\Gamma_3/\Gamma \times B$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
seen		<sup>1</sup> AAIJ	15M LHCb	$p\bar{p}$ at 8 TeV
seen		<sup>2</sup> KHACHATRYAN	15AA CMS	$p\bar{p}$ at 7 TeV
seen		AALTONEN	13 CDF	$p\bar{p}$ at 1.96 TeV
seen		<sup>3</sup> AAIJ	12AV LHCb	$p\bar{p}$ at 7 TeV
seen		AALTONEN	08M CDF	$p\bar{p}$ at 1.96 TeV
seen		ABAZOV	08T D0	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 2.4 \times 10^{-4}$	90	<sup>4</sup> ACKERSTAFF	98O OPAL	$e^+e^- \rightarrow Z$
$< 3.4 \times 10^{-4}$	90	<sup>5</sup> ABREU	97E DLPH	$e^+e^- \rightarrow Z$
$< 8.2 \times 10^{-5}$	90	<sup>6</sup> BARATE	97H ALEP	$e^+e^- \rightarrow Z$
$< 2.0 \times 10^{-5}$	95	<sup>7</sup> ABE	96R CDF	$p\bar{p}$ 1.8 TeV

<sup>1</sup> AAIJ 15M reports a measurement of  $B(B_c^+ \rightarrow J/\psi \pi^+) / B(B^+ \rightarrow J/\psi K^+) \cdot f_c/f_u = (0.683 \pm 0.018 \pm 0.009)\%$  at  $p_T(B) < 20$  GeV and  $2.0 < y(B) < 4.5$ .<sup>2</sup> KHACHATRYAN 15AA reports a measurement of  $B(B_c^+ \rightarrow J/\psi \pi^+) / B(B^+ \rightarrow J/\psi K^+) \cdot f_c/f_u = (0.48 \pm 0.05 \pm 0.03 \pm 0.05)\%$ , at  $p_T > 15$  GeV and  $|\eta(B)| < 1.6$ .<sup>3</sup> AAIJ 12AV reports a measurement of  $B(B_c^+ \rightarrow J/\psi \pi^+)/B(B^+ \rightarrow J/\psi K^+) \cdot f_c/f_u = (0.68 \pm 0.10 \pm 0.03 \pm 0.05)\%$  at  $p_T(B) > 4$  GeV and  $2.5 < \eta(B) < 4.5$ .<sup>4</sup> ACKERSTAFF 98O reports  $B(Z \rightarrow B_c X)/B(Z \rightarrow q\bar{q}) \times B(B_c \rightarrow J/\psi(1S) \pi^+) < 1.06 \times 10^{-4}$  at 90%CL. We rescale to our PDG 98 values of  $B(Z \rightarrow b\bar{b})$ .<sup>5</sup> ABREU 97E value listed is for an assumed  $\tau_{B_c} = 0.4$  ps and improves to  $2.7 \times 10^{-4}$  for  $\tau_{B_c} = 1.4$  ps.<sup>6</sup> BARATE 97H reports  $B(Z \rightarrow B_c X)/B(Z \rightarrow q\bar{q}) \times B(B_c \rightarrow J/\psi(1S) \pi) < 3.6 \times 10^{-5}$  at 90%CL. We rescale to our PDG 96 values of  $B(Z \rightarrow b\bar{b})$ .<sup>7</sup> ABE 96R reports  $B(b \rightarrow B_c X)/B(b \rightarrow B^+ X) \times B(B_c^+ \rightarrow J/\psi(1S) \pi^+)/B(B^+ \rightarrow J/\psi(1S) K^+) < 0.053$  at 95%CL for  $\tau_{B_c} = 0.8$  ps. It changes from 0.15 to 0.04 for  $0.17 \text{ ps} < \tau_{B_c} < 1.6$  ps. We rescale to our PDG 96 values of  $B(b \rightarrow B^+) = 0.378 \pm 0.022$  and  $B(B^+ \rightarrow J/\psi(1S) K^+) = 0.00101 \pm 0.00014$ . $\Gamma(J/\psi(1S) \pi^+)/\Gamma(J/\psi(1S) \mu^+ \nu_\mu)$  $\Gamma_3/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>(4.69 \pm 0.28 \pm 0.46) \times 10^{-2}</math></b>	<sup>1</sup> AAIJ	14W LHCb	$p\bar{p}$ at 7 TeV

<sup>1</sup> AAIJ 14W reports also a measurement  $B(B_c^+ \rightarrow J/\psi \pi^+) / B(B_c^+ \rightarrow J/\psi \mu^+ \nu_\mu) = 0.271 \pm 0.016 \pm 0.016$  in the region  $m_{J/\psi \mu^+} > 5.3$  GeV.

# Meson Particle Listings

 $B_c^+$ 

$\Gamma(J/\psi(1S)K^+)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_4/\Gamma_3$
VALUE	CL%
$0.069 \pm 0.019 \pm 0.005$	50
DOCUMENT ID	TECN COMMENT
AAIJ	13BY LHCb $pp$ at 7 TeV

$\Gamma(J/\psi(1S)\pi^+\pi^+\pi^-)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_5/\Gamma \times B$
VALUE	CL%
seen	
DOCUMENT ID	TECN COMMENT
AAIJ	12Y LHCb $pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$< 5.7 \times 10^{-4}$	90
1 ABREU 97E DLPH $e^+e^- \rightarrow Z$	
1 ABREU 97E value listed is independent of $0.4 \text{ ps} < \tau_{B_c} < 1.4 \text{ ps}$ .	

$\Gamma(J/\psi(1S)\pi^+\pi^+\pi^-)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_5/\Gamma_3$
VALUE	CL%
$2.4 \pm 0.4$ OUR AVERAGE	
2.55 $\pm 0.80 \pm 0.33^{+0.04}_{-0.01}$	
DOCUMENT ID	TECN COMMENT
15AA CMS	$pp$ at 7 TeV
2.41 $\pm 0.30 \pm 0.33$	
AAIJ	12Y LHCb $pp$ at 7 TeV

$\Gamma(J/\psi(1S)a_1(1260))/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_6/\Gamma \times B$
VALUE	CL%
$< 1.2 \times 10^{-3}$	90
DOCUMENT ID	TECN COMMENT
1 ACKERSTAFF 98O OPAL $e^+e^- \rightarrow Z$	
1 ACKERSTAFF 98O reports $B(Z \rightarrow B_c X)/B(Z \rightarrow qq) \times B(B_c \rightarrow J/\psi(1S)a_1(1260)) < 5.29 \times 10^{-4}$ at 90%CL. We rescale to our PDG 98 values of $B(Z \rightarrow b\bar{b})$ .	

$\Gamma(J/\psi(1S)K^+K^-\pi^+)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_7/\Gamma \times B$
VALUE	CL%
seen	
DOCUMENT ID	TECN COMMENT
1 AAIJ	13CA LHCb $pp$ at 7, 8 TeV
1 A signal yield of $78 \pm 14$ decays is reported with a significance of 6.2 standard deviations using an integrated luminosity of $3 \text{ fb}^{-1}$ data.	

$\Gamma(J/\psi(1S)K^+K^-\pi^+)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_7/\Gamma_3$
VALUE	CL%
$0.53 \pm 0.10 \pm 0.05$	
DOCUMENT ID	TECN COMMENT
1 AAIJ	13CA LHCb $pp$ at 7, 8 TeV
1 A signal yield of $78 \pm 14$ decays is reported with a significance of 6.2 standard deviations using an integrated luminosity of $3 \text{ fb}^{-1}$ data.	

$\Gamma(J/\psi(1S)\pi^+\pi^+\pi^-\pi^-)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_8/\Gamma_3$
VALUE	CL%
$1.74 \pm 0.44 \pm 0.24$	
DOCUMENT ID	TECN COMMENT
1 AAIJ	14P LHCb $pp$ at 7, 8 TeV
1 A signal yield of $32 \pm 8$ decays is reported with a significance of 4.5 standard deviations.	

$\Gamma(\psi(2S)\pi^+)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_9/\Gamma_3$
VALUE	CL%
$0.268 \pm 0.032 \pm 0.007 \pm 0.006$	
DOCUMENT ID	TECN COMMENT
1 AAIJ	15AY LHCb $pp$ at 7, 8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.250 \pm 0.068 \pm 0.014 \pm 0.006$	
1 AAIJ	13AMLHCb Repl. by AAIJ 15AY
1 The last uncertainty is due to the uncertainty of the $B(\psi(2S) \rightarrow \mu^+\mu^-)/B(J/\psi \rightarrow \mu^+\mu^-)$ ratio measurement.	

$\Gamma(J/\psi(1S)D_s^+)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_{10}/\Gamma_3$
VALUE	CL%
$3.1 \pm 0.5$ OUR AVERAGE	
3.8 $\pm 1.1 \pm 0.4$	
AAD	16H ATLS $pp$ at 7, 8 TeV
2.90 $\pm 0.57 \pm 0.24$	
AAIJ	13As LHCb $pp$ at 7, 8 TeV

$\Gamma(J/\psi(1S)D_s^{*+})/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_{11}/\Gamma_3$
VALUE	CL%
$10.4 \pm 3.1 \pm 1.6$	
DOCUMENT ID	TECN COMMENT
AAD	16H ATLS $pp$ at 7, 8 TeV

$\Gamma(J/\psi(1S)D_s^{*+})/\Gamma(J/\psi(1S)D_s^+)$	$\Gamma_{11}/\Gamma_{10}$
VALUE	CL%
$2.5 \pm 0.5$ OUR AVERAGE	
2.8 $^{+1.2}_{-0.8} \pm 0.3$	
AAD	16H ATLS $pp$ at 7, 8 TeV
2.37 $\pm 0.56 \pm 0.10$	
AAIJ	13As LHCb $pp$ at 7, 8 TeV

$\Gamma(J/\psi(1S)\rho\bar{\rho}\pi^+)/\Gamma(J/\psi(1S)\pi^+)$	$\Gamma_{12}/\Gamma_3$
VALUE	CL%
$0.143 \pm 0.041_{-0.036}$	
DOCUMENT ID	TECN COMMENT
AAIJ	14AQ LHCb $pp$ at 7, 8 TeV

$\Gamma(D^*(2010)+\bar{D}^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{13}/\Gamma \times B$
VALUE	CL%
$< 6.2 \times 10^{-3}$	90
DOCUMENT ID	TECN COMMENT
1 BARATE 98Q ALEP $e^+e^- \rightarrow Z$	
1 BARATE 98Q reports $B(Z \rightarrow B_c X) \times B(B_c \rightarrow D^*(2010) + \bar{D}^0) < 1.9 \times 10^{-3}$ at 90%CL. We rescale to our PDG 98 values of $B(Z \rightarrow b\bar{b})$ .	

$\Gamma(D^+K^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{14}/\Gamma \times B$
VALUE (units $10^{-6}$ )	CL%
$< 0.20$	90
DOCUMENT ID	TECN COMMENT
1 AAIJ	13R LHCb $pp$ at 7 TeV
1 AAIJ 13R reports $[\Gamma(B_c^+ \rightarrow D^+K^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)] / [B(\bar{B} \rightarrow B^+)] < 0.5 \times 10^{-6}$ which we multiply by our best value $B(\bar{B} \rightarrow B^+) = 40.4 \times 10^{-2}$ .	

$\Gamma(D^+\bar{K}^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{15}/\Gamma \times B$
VALUE (units $10^{-6}$ )	CL%
$< 0.16$	90
DOCUMENT ID	TECN COMMENT
1 AAIJ	13R LHCb $pp$ at 7 TeV
1 AAIJ 13R reports $[\Gamma(B_c^+ \rightarrow D^+\bar{K}^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)] / [B(\bar{B} \rightarrow B^+)] < 0.4 \times 10^{-6}$ which we multiply by our best value $B(\bar{B} \rightarrow B^+) = 40.4 \times 10^{-2}$ .	

$\Gamma(D_s^+K^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{16}/\Gamma \times B$
VALUE (units $10^{-6}$ )	CL%
$< 0.28$	90
DOCUMENT ID	TECN COMMENT
1 AAIJ	13R LHCb $pp$ at 7 TeV
1 AAIJ 13R reports $[\Gamma(B_c^+ \rightarrow D_s^+K^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)] / [B(\bar{B} \rightarrow B^+)] < 0.7 \times 10^{-6}$ which we multiply by our best value $B(\bar{B} \rightarrow B^+) = 40.4 \times 10^{-2}$ .	

$\Gamma(D_s^+\bar{K}^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{17}/\Gamma \times B$
VALUE (units $10^{-6}$ )	CL%
$< 0.4$	90
DOCUMENT ID	TECN COMMENT
1 AAIJ	13R LHCb $pp$ at 7 TeV
1 AAIJ 13R reports $[\Gamma(B_c^+ \rightarrow D_s^+\bar{K}^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)] / [B(\bar{B} \rightarrow B^+)] < 1.1 \times 10^{-6}$ which we multiply by our best value $B(\bar{B} \rightarrow B^+) = 40.4 \times 10^{-2}$ .	

$\Gamma(D_s^+\phi)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{18}/\Gamma \times B$
VALUE (units $10^{-6}$ )	CL%
$< 0.32$	90
DOCUMENT ID	TECN COMMENT
1 AAIJ	13R LHCb $pp$ at 7 TeV
1 AAIJ 13R reports $[\Gamma(B_c^+ \rightarrow D_s^+\phi)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)] / [B(\bar{B} \rightarrow B^+)] < 0.8 \times 10^{-6}$ which we multiply by our best value $B(\bar{B} \rightarrow B^+) = 40.4 \times 10^{-2}$ .	

$\Gamma(K^+K^0)/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{19}/\Gamma \times B$
VALUE	CL%
$< 4.6 \times 10^{-7}$	90
DOCUMENT ID	TECN COMMENT
1 AAIJ	13Bs LHCb $pp$ at 7 TeV
1 Derived from $\Gamma(K^+K^0)/\Gamma \times B(\bar{B} \rightarrow B_c) / (B(B^+ \rightarrow K^0\pi^+)B(\bar{B} \rightarrow B^+)) < 5.8\%$ at 90% CL using normalization mode $B(B^+ \rightarrow K^0\pi^+) = (23.97 \pm 0.53 \pm 0.71) \times 10^{-6}$ and assuming a $B$ production ratio $f(\bar{B} \rightarrow B_u^+) = 0.33$ .	

$\Gamma(B_s^+\pi^+/B(\bar{B} \rightarrow B_s))/\Gamma_{\text{total}} \times B(\bar{B} \rightarrow B_c)$	$\Gamma_{20}/\Gamma \times B$
VALUE (units $10^{-3}$ )	CL%
$2.37 \pm 0.31 \pm 0.11^{+0.17}_{-0.13}$	
DOCUMENT ID	TECN COMMENT
1 AAIJ	13Bu LHCb $pp$ at 7, 8 TeV
1 The last uncertainty is due to the uncertainty of the $B_c^+$ lifetime measurement.	

## POLARIZATION IN $B_c^+$ DECAY

In decays involving two vector mesons, one can distinguish among the states in which meson polarizations are both longitudinal (L) or both are transverse and parallel (||) or perpendicular ( $\perp$ ) to each other with the parameters  $\Gamma_L/\Gamma$ ,  $\Gamma_{\perp}/\Gamma$ , and the relative phases  $\phi_{||}$  and  $\phi_{\perp}$ . See the definitions in the note on “Polarization in  $B$  Decays” review in the  $B^0$  Particle Listings.

$\Gamma_L/\Gamma$ in $B_c^+ \rightarrow J/\psi D_s^{*+}$	DOCUMENT ID	TECN	COMMENT
VALUE			
$0.54 \pm 0.15$ OUR AVERAGE			
0.62 $\pm 0.24$	1 AAD	16H ATLS	$pp$ at 7, 8 TeV
0.48 $\pm 0.20$	2 AAIJ	13As LHCb	$pp$ at 7, 8 TeV
1 AAD 16H measures $1 - \Gamma_L/\Gamma = 0.38 \pm 0.24$ .			
2 AAIJ 13As measures $1 - \Gamma_L/\Gamma = 0.52 \pm 0.20$ .			

## $B_c^+$ REFERENCES

AAD	16H	EPJ C76 4	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAU	15AY	PR D92 072007	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	15G	PL B742 29	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAU	15M	PRL 114 132001	R. Aaij <i>et al.</i>	(LHCb Collab.)
KHACHATRY...	15AA	JHEP 1501 063	V. Khachatryan <i>et al.</i>	(CMS Collab.)
AAIJ	14AQ	PRL 113 152003	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14G	EPJ C74 2839	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14P	JHEP 1405 148	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14W	PR D90 032009	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13AM	PR D87 071103	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13AS	PR D87 112012	R. Aaij <i>et al.</i>	(LHCb Collab.)
Also		PR D89 019901 (err.)	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13BS	PL B726 646	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13BU	PRL 111 181801	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13BY	JHEP 1309 075	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13CA	JHEP 1311 094	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13R	JHEP 1302 043	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	13	PR D87 011101	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AAIJ	12AV	PRL 109 232001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	12Y	PRL 108 251802	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	09M	PRL 102 092001	V.M. Abazov <i>et al.</i>	(DO Collab.)
AALTONEN	08H	PRL 100 182002	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	08T	PRL 101 012001	V.M. Abazov <i>et al.</i>	(DO Collab.)
ABULENCIA	06C	PRL 96 082002	A. Abulencia <i>et al.</i>	(CDF Collab.)
ABULENCIA	06O	PRL 97 012002	A. Abulencia <i>et al.</i>	(CDF Collab.)
ABE	98M	PRL 81 2432	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PR D58 112004	F. Abe <i>et al.</i>	(CDF Collab.)
ACKERSTAFF	98O	PL B420 157	K. Ackersstaff <i>et al.</i>	(OPAL Collab.)
BARATE	98Q	EPJ C4 387	R. Barate <i>et al.</i>	(ALEPH Collab.)
PDG	98	EPJ C3 1	C. Caso <i>et al.</i>	(PDG Collab.)
ABREU	97E	PL B398 207	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BARATE	97H	PL B402 213	R. Barate <i>et al.</i>	(ALEPH Collab.)
ABE	96R	PRL 77 5176	F. Abe <i>et al.</i>	(CDF Collab.)
PDG	96	PR D54 1	R. M. Barnett <i>et al.</i>	(PDG Collab.)

$B_c(2S)^\pm$ , Heavy Quarkonium Spectroscopy $B_c(2S)^\pm$ 

$$I(J^P) = 0(0^-)$$

OMITTED FROM SUMMARY TABLE

Quantum numbers neither measured nor confirmed.

 $B_c(2S)^\pm$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6842 \pm 4 \pm 5</math></b>	57	<sup>1</sup> AAD	14AQ ATLS	$pp$ at 7, 8 TeV

<sup>1</sup> Observed in the decay mode  $B_c(2S)^+ \rightarrow B_c^+ \pi^+ \pi^-$  ( $B_c^+ \rightarrow J/\psi \pi^+$ ) with 5.2 standard deviations significance.

 $B_c(2S)^\pm$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $B_c^+ \pi^+ \pi^-$	seen

 $B_c(2S)^\pm$  BRANCHING RATIOS

$\Gamma(B_c^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	57	<sup>1</sup> AAD	14AQ ATLS	$pp$ at 7, 8 TeV

<sup>1</sup> Observed with 5.2 standard deviations significance.

 $B_c(2S)^\pm$  REFERENCES

AAD      14AQ PRL 113 212004      G. Aad et al.      (ATLAS Collab.)

## SPECTROSCOPY OF MESONS CONTAINING TWO HEAVY QUARKS

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A golden age for heavy quarkonium physics dawned at the turn of this century, initiated by the confluence of exciting advances in quantum chromodynamics (QCD) and an explosion of related experimental activity. The subsequent broad spectrum of breakthroughs, surprises, and continuing puzzles had not been anticipated. In that period, the BESIII program concluded only to give birth to BESIII; the  $B$ -factories and CLEO-c flourished; quarkonium production and polarization measurements at HERA and the Tevatron matured; and heavy-ion collisions at RHIC opened a window on the deconfinement regime. Recently also ATLAS, CMS and LHCb started to contribute to the field. For an extensive presentation of the status of heavy quarkonium physics, the reader is referred to several reviews [1–8]. This note focuses on experimental developments in heavy quarkonium spectroscopy with very few theoretical comments. Some other comments on possible theoretical interpretations of the states not predicted by the quark model are presented in the mini-review on non  $\bar{q}q$ -states.

In this mini-review we display the newly discovered states, where “newly” is interpreted to include the period since 2002. In earlier versions of this write-up the particles were sorted according to an assumed *conventional* or *unconventional* nature

with respect to the quark model. However, since this classification is not always unambiguous, we here follow Ref. [8] and sort the states into three groups, namely states below (*cf.* Table 1), near (*cf.* Table 2) and above (*cf.* Table 3) the lowest open flavor thresholds.

Table 1 lists properties of newly observed heavy quarkonium states located below the lowest open flavor thresholds. Those are expected to be (at least prominently) conventional quarkonia. The  $h_c(1P)$  is the  $^1P_1$  state of charmonium, singlet partner of the long-known  $\chi_{cJ}$  triplet  $^3P_J$ . The  $\eta_c(2S)$  is the first excited state of the pseudoscalar ground state  $\eta_c(1S)$ , lying just below the mass of its vector counterpart,  $\psi(2S)$ .

Although  $\eta_c(2S)$  measurements began to converge towards a mass and a width some time ago, refinements are still in progress. In particular, Belle [16] has revisited its analysis of  $B \rightarrow K\eta_c(2S)$ ,  $\eta_c(2S) \rightarrow KK\pi$  decays with more data and methods that account for interference between the above decay chain, an equivalent one with the  $\eta_c(1S)$  instead, and one with no intermediate resonance. The net effect of this interference is far from trivial; it shifts the apparent mass by  $\sim +10$  MeV and blows up the apparent width by a factor of six. The updated  $\eta_c(2S)$  mass and width are in better accordance with other measurements than the previous treatment [15], which did not include interference. Complementing this measurement in  $B$ -decay, BaBar [17] updated their previous [18]  $\eta_c(2S)$  mass and width measurements in two-photon production, where interference effects, judging from studies of  $\eta_c(1S)$ , appear to be small. In combination, precision on the  $\eta_c(2S)$  mass has improved dramatically. In addition, Belle recently reported a measurement of  $\psi_2(1D)$  which would be a  $J^{PC} = 2^{+-}$  state [23]. Its existence was confirmed with high significance by BESIII [24]. While the negative C-parity is indeed established by the measurement, the assignment of  $J = 2$  was done by matching to the closest quark model state. In the table this state is therefore simply called  $X(3823)$ , according to the PDG name convention.

A new  $\bar{c}b$  state was discovered by the ATLAS Collaboration [28]. They observed an excited  $B_c^\pm$  state, which properties are consistent with expectations for the second  $S$ -wave state of the  $B_c^\pm$  meson,  $B_c^\pm(2S)$ .

The ground state of bottomonium is the  $\eta_b(1S)$ , recently confirmed with a second observation of more than  $5\sigma$  significance at Belle. In addition, in the same experiment strong evidence was collected for  $\eta_b(2S)$  [32], but it still needs experimental confirmation at the  $5\sigma$  level. The  $\Upsilon(1D)$  is the lowest-lying  $D$ -wave triplet of the  $b\bar{b}$  system. Both the  $h_b(1P)$ , the bottomonium counterpart of  $h_c(1P)$ , and the next excited state,  $h_b(2P)$ , were recently observed by Belle [35], as described further below, in dipion transitions from the  $\Upsilon(10860)$ . We no longer mention a hypothetical  $Y_b(10888)$  state since new analysis of the  $\Upsilon(10860)$  energy range does not show evidence for an additional state with mass shifted from the  $\Upsilon(10860)$  [111]. After the mass of the  $\eta_b(1S)$  was shifted upwards by about 10 MeV based on the new Belle measurements [32,33], all states

# Meson Particle Listings

## Heavy Quarkonium Spectroscopy

mentioned in this paragraph fit into their respective spectroscopies roughly where expected. Their exact masses, production mechanisms, and decay modes provide guidance to their descriptions within QCD.

them have been confirmed experimentally as indicated in the last column of the tables. With the possible exception of the tensor state located at 3930 MeV, neither can unambiguously be assigned a place in the hierarchy of

**Table 1:** New states below the open flavor thresholds in the  $c\bar{c}$ ,  $b\bar{c}$ , and  $b\bar{b}$  regions, ordered by mass. Masses  $m$  and widths  $\Gamma$  represent the PDG16 weighted averages. Ellipses (...) in the Process column indicate inclusively selected event topologies; *i.e.*, additional particles not required by the Experiments to be present. A question mark (?) indicates an unmeasured value. For each Experiment a citation is given, as well as the statistical significance ( $\#\sigma$ ), or “(np)” for “not provided”. The Year column gives the date of the first measurement cited. The Status column indicates that the state has been observed by at most one (NC!=needs confirmation) or at least two independent experiments with significance of  $>5\sigma$  (OK).

State	$m$ (MeV)	$\Gamma$ (MeV)	$J^{PC}$	Process (mode)	Experiment ( $\#\sigma$ )	Year	Status
$h_c(1P)$	$3525.38 \pm 0.11$	$0.7 \pm 0.35$	$1^{+-}$	$\psi(2S) \rightarrow \pi^0(\gamma\eta_c(1S))$ $\psi(2S) \rightarrow \pi^0(\gamma\dots)$ $p\bar{p} \rightarrow (\gamma\eta_c) \rightarrow (\gamma\gamma\gamma)$ $\psi(2S) \rightarrow \pi^0(\gamma\eta_c(1S))$	CLEO [9–11] (13.2) CLEO [9–11] (10), BES [12] (19) E835 [13] (3.1) BESIII [14] (np)	2004	OK
$\eta_c(2S)$	$3639.2 \pm 1.2$	$11.3^{+3.2}_{-2.9}$	$0^{-+}$	$B \rightarrow K(K_S^0 K^- \pi^+)$ $e^+e^- \rightarrow e^+e^-(K_S^0 K^- \pi^+)$ $e^+e^- \rightarrow J/\psi(\dots)$	Belle [15,16] (6.0) BaBar [17,18] (7.8), CLEO [19] (6.5), Belle [20] (6) BaBar [21] (np), Belle [22] (8.1)	2002	OK
$X(3823)$	$3822.5 \pm 1.2$	$< 16$	$?^{-}$	$B \rightarrow K(\gamma\chi_{c1})$ $e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}\gamma$	Belle [23](3.8) BESIII [24] (6.2)	2013	NC!
$B_c^+$	$6277 \pm 6$	?	$0^-$	$\bar{p}p \rightarrow (\pi^+ J/\psi)\dots$	CDF [25,26] (8.0), D0 [27] (5.2)	2007	OK
$B_c^+(2S)$	$6842 \pm 6$	?	$0^-$	$pp \rightarrow (B_c^+ \pi^+ \pi^-)\dots$	ATLAS [28] (5.2)	2014	NC!
$\eta_b(1S)$	$9399.2 \pm 1.9$	$9.8^{+4.4}_{-3.6}$	$0^{-+}$	$\Upsilon(3S) \rightarrow \gamma(\dots)$ $\Upsilon(2S) \rightarrow \gamma(\dots)$ $h_b(1P, 2P) \rightarrow \gamma(\dots)$ $\Upsilon(4S) \rightarrow \eta h_b(1P)$ $\Upsilon(10860) \rightarrow \pi^+\pi^-\gamma(\dots)$	BaBar [29] (10), CLEO [30] (4.0) BaBar [31] (3.0) Belle [32](14) Belle [33](9) Belle [34] (14)	2008	OK
$h_b(1P)$	$9899.3 \pm 0.7$	?	$1^{+-}$	$\Upsilon(10860) \rightarrow \pi^+\pi^-(\dots)$ $\Upsilon(3S) \rightarrow \pi^0(\dots)$ $\Upsilon(4S) \rightarrow \eta h_b(1P)$	Belle [35,34] (5.5) BaBar [36] (3.0) Belle [33] (11)	2011	NC!
$\eta_b(2S)$	$9999.0^{+4.5}_{-4.0}$	$< 24$	$0^{-+}$	$h_b(2P) \rightarrow \gamma(\dots)$	Belle [32](4.2)	2012	NC!
$\Upsilon(1^3D_2)$	$10163.7 \pm 1.4$	?	$2^{--}$	$\Upsilon(3S) \rightarrow \gamma\gamma(\gamma\gamma\Upsilon(1S))$ $\Upsilon(3S) \rightarrow \gamma\gamma(\pi^+\pi^-\Upsilon(1S))$ $\Upsilon(10860) \rightarrow \pi^+\pi^-(\dots)$	CLEO [37] (10.2) BaBar [38] (5.8) Belle [35] (2.4)	2004	OK
$h_b(2P)$	$10259.8^{+1.5}_{-1.2}$	?	$1^{+-}$	$\Upsilon(10860) \rightarrow \pi^+\pi^-(\dots)$	Belle [35,34] (11.2)	2011	NC!
$\chi_{bJ}(3P)$	$10512.1 \pm 2.3$	?	$?^{2+}$	$pp \rightarrow (\gamma\mu^+\mu^-)\dots$	ATLAS [39] ( $>6$ ), D0 [40] (3.6) LHCb [41] (6.9)	2011	OK

There is a large number of newly discovered states both near and above the lowest open flavor thresholds. They are displayed in Table 2 and Table 3, respectively\*; notice that just a few of

\* For consistency with the literature, we preserve the use of  $X$ ,  $Y$  and  $Z$ , contrary to the practice of the PDG, which exclusively uses  $X$  for states with undetermined quantum numbers.

charmonia or bottomonia. However, besides the charged states, none has a universally accepted unconventional origin either. The  $X(3872)$  is widely studied, yet its interpretation demands additional experimental attention: after the quantum numbers were fixed at LHCb [59,60], the next experimental challenge will be a measurement of its line shape. The state originally

dubbed  $Z(3930)$  is now regarded by many as the first observed  $2P$  state of  $\chi_{cJ}$ , the  $\chi_{c2}(2P)$ . Another state was discovered at 3915 MeV [75] and from a subsequent measurement its quantum numbers were determined to be  $J^{PC} = 0^{++}$  [77] suggesting it to be the  $\chi_{c0}(2P)$  quark model state, but this interpretation is not generally accepted [114,115]. In addition, it was pointed out in Ref. [116] that if the assumption of a helicity-2 dominance is abandoned and instead one allows for a sizable helicity-0 component, a  $J^{PC} = 2^{++}$  assignment is possible. This could imply that the state at 3930 MeV is actually identical to the one at 3915 MeV—but to explain the large helicity-0 component a sizable portion of non- $\bar{q}q$  is necessary [116]. Because of this analysis the name of the state was changed back from  $\chi_{c0}(2P)$  to  $X(3915)$ .

Based on a full amplitude analysis of the  $B^0 \rightarrow K^+ \pi^- \psi(2S)$  decays, Belle determined the spin-parity of the  $Z(4430)^{\pm**}$  to be  $J^P = 1^+$  [105]. Very recently this state as well as its quantum numbers were confirmed at LHCb [107] with much higher statistics. Improved values for mass and width from LHCb are consistent with earlier measurements; our new average is in Table 3; the experiment even reports a resonant behavior of the  $Z(4430)^{\pm}$  amplitude. This state as well as  $Z(4050)^{\pm}$  and  $Z(4250)^{\pm}$  seen in  $\pi^{\pm}\chi_{c1}$  is, however, not confirmed (nor excluded) by BaBar (see [106] for the  $Z(4430)$  and [83] for the  $Z(4050)^{\pm}$  and  $Z(4250)^{\pm}$ ). Belle observes signals of significances  $5.0\sigma$ ,  $5.0\sigma$ , and  $6.4\sigma$  for  $Z_1(4050)^+$ ,  $Z_2(4250)^+$ , and  $Z(4430)^+$ , respectively, whereas BABAR reports  $1.1\sigma$ ,  $2.0\sigma$ , and  $2.4\sigma$  effects, setting upper limits on product branching

**Table 2:** As in Table 1, but for new states near the first open flavor thresholds in the  $c\bar{c}$  and  $b\bar{b}$  regions, ordered by mass. For  $X(3872)$ , the values given are based only upon decays to  $\pi^+\pi^-J/\psi$ . Updated from [7] with kind permission, copyright (2011), Springer, and [8] with kind permission from the authors.

State	$m$ (MeV)	$\Gamma$ (MeV)	$J^{PC}$	Process (mode)	Experiment ( $\# \sigma$ )	Year	Status
$X(3872)$	$3871.68 \pm 0.17$	$< 1.2$	$1^{++}$	$B \rightarrow K(\pi^+\pi^-J/\psi)$ $p\bar{p} \rightarrow (\pi^+\pi^-J/\psi) + \dots$ $B \rightarrow K(\omega J/\psi)$ $B \rightarrow K(D^{*0}\bar{D}^0)$ $B \rightarrow K(\gamma J/\psi)$  $B \rightarrow K(\gamma\psi(2S))$	Belle [42,43] (10.3), BaBar [44] (8.6) CDF [45–47] (np), D0 [48] (5.2) Belle [49] (4.3), BaBar [50] (4.9) Belle [51,52] (6.4), BaBar [53] (4.9) Belle [54] (4.0), BaBar [55,56] (3.6), LHCb [57] ( $>10$ ) BaBar [56] (3.5), Belle [54] (0.4), LHCb [57] (4.4)	2003	OK
$Z_c(3900)$	$3891.2 \pm 3.3$	$40 \pm 8$	$1^{+-}$	$pp \rightarrow (\pi^+\pi^-J/\psi) + \dots$ $Y(4260) \rightarrow \pi^-(\pi^+J/\psi)$  $Y(4260) \rightarrow \pi^0(\pi^0J/\psi)$	BESIII [61] ( $> 8$ ), Belle [62] (5.2) CLEO data [63] ( $> 5$ ) BESIII [64] (10.4) CLEO data [63] (3.5)	2013	OK
$Z_c(4020)$	$4022.9 \pm 2.8$	$7.9 \pm 3.7$	$1^{+-}$	$Y(4260) \rightarrow \pi^-(D\bar{D}^*)^+$ $Y(4260) \rightarrow \pi^0(D\bar{D}^*)^0$ $Y(4260, 4360) \rightarrow \pi^-(\pi^+h_c)$ $Y(4260, 4360) \rightarrow \pi^0(\pi^0h_c)$ $Y(4260) \rightarrow \pi^-(D^*\bar{D}^*)^+$ $Y(4260) \rightarrow \pi^0(D^*\bar{D}^*)^0$	BESIII [65] (18) BESIII [66] ( $> 10$ ) BESIII [67] (8.9) BESIII [68] ( $> 5$ ) BESIII [69] (10) BESIII [70] (5.9)	2013	NC!
$Z_b(10610)$	$10607.2 \pm 2.0$	$18.4 \pm 2.4$	$1^{+-}$	$\Upsilon(10860) \rightarrow \pi^-(\pi^+\Upsilon(1S, 2S, 3S))$ $\Upsilon(10860) \rightarrow \pi^-(\pi^+h_b(1P, 2P))$ $\Upsilon(10860) \rightarrow \pi^0(\pi^0\Upsilon(1S, 2S, 3S))$ $\Upsilon(10860) \rightarrow \pi^-(B\bar{B}^*)^+$	Belle [71] ( $> 10$ ) [72] Belle [71] (16) Belle [73] (6.5) Belle [74] ( $> 8$ )	2011	NC!
$Z_b(10650)$	$10652.2 \pm 1.5$	$11.5 \pm 2.2$	$1^{+-}$	$\Upsilon(10860) \rightarrow \pi^-(\pi^+\Upsilon(1S, 2S, 3S))$ $\Upsilon(10860) \rightarrow \pi^-(\pi^+h_b(1P, 2P))$ $\Upsilon(10860) \rightarrow \pi^-(B^*\bar{B}^*)^+$	Belle [71] ( $> 10$ ) Belle [71] (16) Belle [74] (6.8)	2011	OK

The  $Y(4260)$  and  $Y(4360)$  are vector states decaying to  $\pi^+\pi^-J/\psi$  and  $\pi^+\pi^-\psi(2S)$ , respectively, yet, unlike most conventional vector charmonia, do not correspond to enhancements in the  $e^+e^-$  hadronic cross section. Another interesting question is whether a heavier  $\pi^+\pi^-\psi(2S)$  state, the  $Y(4660)$ , discovered by Belle [101,102] and confirmed by BaBar [100], is identical to the  $\Lambda_c^+\Lambda_c^-$  state with close parameters observed by Belle using initial-state radiation [108].

fractions that are not inconsistent with Belle’s and LHCb’s measured rates. For the  $Z_1(4050)^+$  and  $Z_2(4250)^+$  states the situation remains unresolved.

\*\* There are currently various candidates for isotriplet states in the spectrum. For some of them both charged states are already established and sometimes there is also evidence for the neutral partner. We still chose to put the charge as superscript since it is an explicit marker of the exotic nature of the states.

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**Table 3:** As in Table 1, but for new states above the first open flavor thresholds in the  $c\bar{c}$  and  $b\bar{b}$  regions, ordered by mass.  $X(3945)$  and  $Y(3940)$  have been subsumed under  $X(3940)$  due to compatible properties. The  $\chi_{c0}(3915)$  is now changed back to  $X(3915)$  as explained in the main text. The state known as  $Z(3930)$  appears as the  $\chi_{c2}(2P)$  in Table 1. In some cases experiment still allows two  $J^{PC}$  values, in which case both appear. See also the reviews in [1–8].

State	$m$ (MeV)	$\Gamma$ (MeV)	$J^{PC}$	Process (mode)	Experiment ( $\# \sigma$ )	Year	Status
$X(3915)$	$3917.4 \pm 2.7$	$28^{+10}_{-9}$	$0/2^{++}$	$B \rightarrow K(\omega J/\psi)$ $e^+e^- \rightarrow e^+e^- \omega J/\psi$	Belle [75] (8.1), BaBar [50] (np) Belle [76] (7.7), BaBar [77] (19)	2004	OK
$\chi_{c2}(2P)$	$3927.2 \pm 2.6$	$24 \pm 6$	$2^{++}$	$e^+e^- \rightarrow e^+e^- (D\bar{D})$	Belle [78] (5.3), BaBar [79]	2005	OK
$X(3940)$	$3942^{+9}_{-8}$	$37^{+27}_{-17}$	$?^{?+}$	$e^+e^- \rightarrow J/\psi (D\bar{D}^*)$ $e^+e^- \rightarrow J/\psi (\dots)$	Belle [80] (6.0) Belle [22] (5.0)	2007	NC!
$Y(4008)$	$4008^{+121}_{-49}$	$226 \pm 97$	$1^{--}$	$e^+e^- \rightarrow \gamma(\pi^+\pi^- J/\psi)$	Belle [81] (7.4)	2007	NC!
$Z_1(4050)^+$	$4051^{+24}_{-43}$	$82^{+51}_{-55}$	$?$	$B \rightarrow K(\pi^+\chi_{c1}(1P))$	Belle [82] (5.0), BaBar [83] (1.1)	2008	NC!
$Y(4140)$	$4145.8 \pm 2.6$	$18 \pm 8$	$?^{?+}$	$B^+ \rightarrow K^+(\phi J/\psi)$	CDF [84,85] (5.0) D0 [86] (3.1), CMS [87] (>5) Belle [88] (1.9), LHCb [89] (1.4), BaBar [90]	2009	NC!
$X(4160)$	$4156^{+29}_{-25}$	$139^{+113}_{-65}$	$?^{?+}$	$e^+e^- \rightarrow e^+e^- (\phi J/\psi)$ $e^+e^- \rightarrow J/\psi (D\bar{D}^*)$	Belle [91] (3.2) Belle [80] (5.5)	2009 2007	NC! NC!
$Z_c(4200)^+$	$4196^{+35}_{-32}$	$370^{+99}_{-149}$	$1^+$	$\bar{B}^0 \rightarrow K^-(J/\psi \pi^+)$	Belle [92] (6.2)	2014	NC!
$Z_2(4250)^+$	$4248^{+185}_{-45}$	$177^{+321}_{-72}$	$?$	$B \rightarrow K(\pi^+\chi_{c1}(1P))$	Belle [82] (5.0), BaBar [83] (2.0)	2008	NC!
$Y(4260)$	$4263^{+8}_{-9}$	$95 \pm 14$	$1^{--}$	$e^+e^- \rightarrow \gamma(\pi^+\pi^- J/\psi)$	BaBar [93,94] (8.0) CLEO [95] (5.4), Belle [81] (15) CLEO [96] (11) CLEO [96] (5.1)	2005	OK
$Y(4274)$	$4293 \pm 20$	$35 \pm 16$	$?^{?+}$	$e^+e^- \rightarrow (\pi^+\pi^- J/\psi)$ $e^+e^- \rightarrow (\pi^0\pi^0 J/\psi)$ $e^+e^- \rightarrow (f_0(980)J/\psi)$ $e^+e^- \rightarrow (\pi^- Z_c(3900)^+)$ $e^+e^- \rightarrow (\gamma X(3872))$ $B^+ \rightarrow K^+(\phi J/\psi)$	BESIII [61] (8), Belle [62] (5.2) BESIII [98] (5.3) CDF [85] (3.1), LHCb [89] (1.0), CMS [87] (>3), D0 [86] (np) Belle [91] (3.2)	2011	NC!
$X(4350)$	$4350.6^{+4.6}_{-5.1}$	$13.3^{+18.4}_{-10.0}$	$0/2^{++}$	$e^+e^- \rightarrow e^+e^- (\phi J/\psi)$	Belle [91] (3.2)	2009	NC!
$Y(4360)$	$4361 \pm 13$	$74 \pm 18$	$1^{--}$	$e^+e^- \rightarrow \gamma(\pi^+\pi^- \psi(2S))$	BaBar [99,100] (np), Belle [101,102] (8.0)	2007	OK
$Z(4430)^+$	$4458 \pm 15$	$166^{+37}_{-32}$	$1^+$	$\bar{B}^0 \rightarrow K^-(\pi^+ \psi(2S))$	Belle [103,104,105] (6.4), BaBar [106] (2.4), LHCb [107] (13.9) Belle [92] (4.0)	2007	OK
$X(4630)$	$4634^{+9}_{-11}$	$92^{+41}_{-32}$	$1^{--}$	$\bar{B}^0 \rightarrow (J/\psi \pi^+) K^-$ $e^+e^- \rightarrow \gamma(\Lambda_c^+ \Lambda_c^-)$	Belle [108] (8.2)	2007	NC!
$Y(4660)$	$4664 \pm 12$	$48 \pm 15$	$1^{--}$	$e^+e^- \rightarrow \gamma(\pi^+\pi^- \psi(2S))$	Belle [101,102] (5.8), BaBar [100] (np)	2007	NC!
$\Upsilon(10860)$	$10876 \pm 11$	$55 \pm 28$	$1^{--}$	$e^+e^- \rightarrow (B_{(s)}^* \bar{B}_{(s)}^* (\pi))$ $e^+e^- \rightarrow (\pi\pi \Upsilon(1S, 2S, 3S))$ $e^+e^- \rightarrow (f_0(980) \Upsilon(1S))$ $e^+e^- \rightarrow (\pi Z_b(10610, 10650))$ $e^+e^- \rightarrow (\eta \Upsilon(1S, 2S))$ $e^+e^- \rightarrow (\pi^+\pi^- \Upsilon(1D))$ $e^+e^- \rightarrow (\pi^+\pi^- h_b(1P, 2P))$	PDG [109] (> 10) Belle [110,71,73,111] (>10) Belle [71,73] (>5) Belle [71,73] (>10) Belle [33] (10) Belle [112] (9) Belle [113] (9)	1985	OK
$\Upsilon(11020)$	$10987.5^{+11.1}_{-3.3}$	$61.0^{+9.2}_{-27.7}$	$1^{--}$	$e^+e^- \rightarrow (B_{(s)}^* \bar{B}_{(s)}^* (\pi))$ $e^+e^- \rightarrow (\pi\pi \Upsilon(1S, 2S, 3S))$ $e^+e^- \rightarrow (\pi^+\pi^- h_b(1P, 2P))$	PDG [109] (> 10) [111] (>10) Belle [113] (9)	1985	OK

In addition to the three  $Z_c^+$  discussed in the previous paragraph, in 2013 two more states named  $Z_c(3900)^+$  and  $Z_c(4020)^+$  were unearthed in the charmonium region. Note that in this write-up as well as the RPP listings we combined  $Z_c(3900)^+$  (seen in  $J/\psi \pi \pi$ ) and  $Z_c(3885)^+$  (seen in  $D\bar{D}^*$ ) as well as  $Z_c(4020)^+$  (seen in  $h_c \pi \pi$ ) and  $Z_c(4025)^+$  (seen in  $D^* \bar{D}^*$ ) into only two states due to their close proximity in mass. In various respects  $Z_c(3900)^+$  and  $Z_c(4020)^+$  seem to be the charmed partners of  $Z_b(10610)^+$  and  $Z_b(10650)^+$  as will be outlined below. Finally, from their study of  $\bar{B}^0 \rightarrow J/\psi K^- \pi^+$

decays Belle reported evidence for one more charged state, dubbed  $Z_c(4200)^+$  [92]. This very analysis gave evidence for the decay mode  $Z(4430) \rightarrow J/\psi \pi$ , which has an order of magnitude lower branching fraction than the discovery mode  $Z(4430) \rightarrow \psi(2S) \pi$ .

The  $Y(4140)$  observed in 2008 by CDF [84,85] was confirmed at D0 and CMS [86,87], however, a second structure related to  $Y(4274)$  could not be established unambiguously. The two states were neither seen in  $B$  decays at Belle [88], LHCb [89]

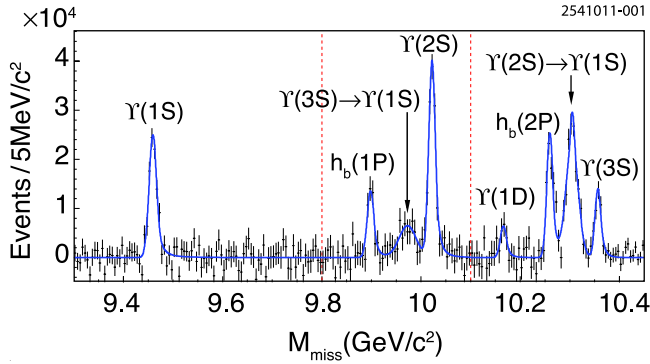
See key on page 601

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and BaBar [90] nor in  $\gamma\gamma$  collisions at Belle [91]. Thus the situation for the  $Y(4140)$  and  $Y(4274)$  is still controversial.

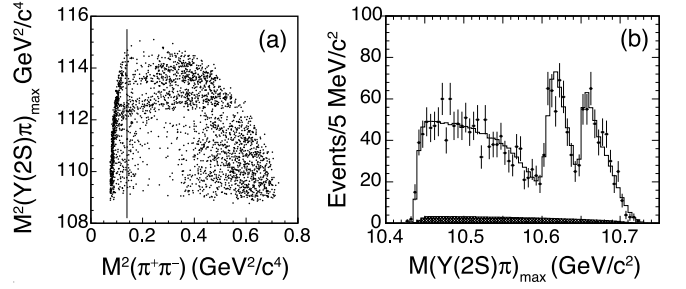
New results on  $\eta_b$ ,  $h_b$ , and  $Z_b^+$  mostly come from Belle [32–35], [71–74], [110–113], all from analyses of  $121.4 \text{ fb}^{-1}$  of  $e^+e^-$  collision data collected near the peak of the  $Y(10860)$  resonance as well as from additional  $25 \text{ fb}^{-1}$  of data collected during the scans of the c.m. energy range 10.63–11.05 GeV. They all appear in the decay chains:  $Y(10860) \rightarrow \pi^- Z_b^+$ ,  $Z_b^+ \rightarrow \pi^+(b\bar{b})$ , and, when the  $b\bar{b}$  forms an  $h_b(1P)$ , frequently decaying as  $h_b(1P) \rightarrow \gamma\eta_b$ .



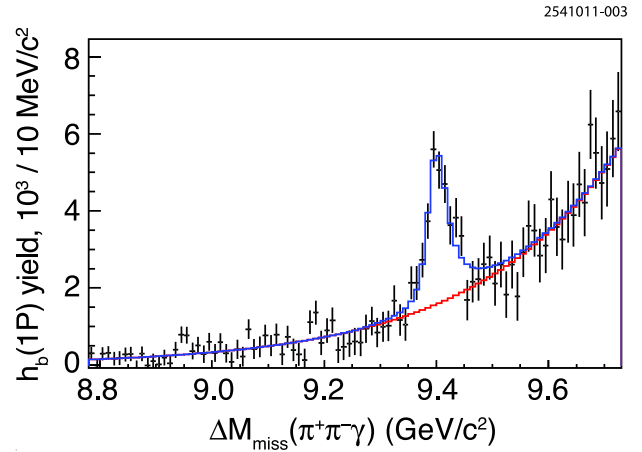
**Figure 1:** From Belle [35], the mass recoiling against  $\pi^+\pi^-$  pairs,  $M_{\text{miss}}$ , in  $e^+e^-$  collision data taken near the peak of the  $Y(10860)$  (points with error bars). The smooth combinatorial and  $K_S^0 \rightarrow \pi^+\pi^-$  background contributions have been subtracted. The fit to the various labeled signal contributions is overlaid (curve). Adapted from [35] with kind permission, copyright (2011) The American Physical Society.

The Belle  $h_b$  discovery analysis [35] selects hadronic events and searches for peaks in the mass recoiling against  $\pi^+\pi^-$  pairs, the spectrum for which, after subtraction of smooth combinatorial and  $K_S^0 \rightarrow \pi^+\pi^-$  backgrounds, appears in Fig. 1. Prominent and unmistakable  $h_b(1P)$  and  $h_b(2P)$  peaks are present. This search was directly inspired by a CLEO result [117], which found the surprisingly copious transitions  $\psi(4160) \rightarrow \pi^+\pi^-h_c(1P)$  and an indication that  $Y(4260) \rightarrow \pi^+\pi^-h_c(1P)$  occurs at a comparable rate as the signature mode,  $Y(4260) \rightarrow \pi^+\pi^-J/\psi$ . The presence of  $Y(nS)$  peaks in Fig. 1 at rates two orders of magnitude larger than expected, along with separate studies with exclusive decays  $Y(nS) \rightarrow \mu^+\mu^-$ , allow precise calibration of the  $\pi^+\pi^-$  recoil mass spectrum and very accurate measurements of  $h_b(1P)$  and  $h_b(2P)$  masses. Both corresponding hyperfine splittings are consistent with zero within an uncertainty of about 1.5 MeV (lowered to  $\pm 1.1 \text{ MeV}$  for  $h_b(1P)$  in Ref. [34]).

Belle soon noticed that, for events in the peaks of Fig. 1, there seemed to be two intermediate charged states nearby. For example, Fig. 2 shows a Dalitz plot for events restricted to the  $Y(2S)$  region of  $\pi^+\pi^-$  recoil mass, with  $Y(2S) \rightarrow \mu^+\mu^-$ . The two bands observed in the maximum of the two  $M[\pi^\pm Y(2S)]^2$  values also appear for  $Y(1S)$ ,  $Y(3S)$ ,  $h_b(1P)$ , and  $h_b(2P)$



**Figure 2:** From Belle [71]  $e^+e^-$  collision data taken near the peak of the  $Y(10860)$  for events with a  $\pi^+\pi^-$ -missing mass consistent with an  $Y(2S) \rightarrow \mu^+\mu^-$ , (a) the maximum of the two possible single  $\pi^\pm$ -missing-mass-squared combinations vs. the  $\pi^+\pi^-$ -mass-squared; and (b) projection of the maximum of the two possible single  $\pi^\pm$ -missing-mass combinations (points with error bars) overlaid with a fit (curve). Events to the left of the vertical line in (a) are excluded from amplitude analysis. The hatched histogram in (b) corresponds to the combinatorial background. The two horizontal stripes in (a) and two peaks in (b) correspond to the two  $Z_b^+$  states. Adapted from [71] with kind permission, copyright (2011) The American Physical Society.



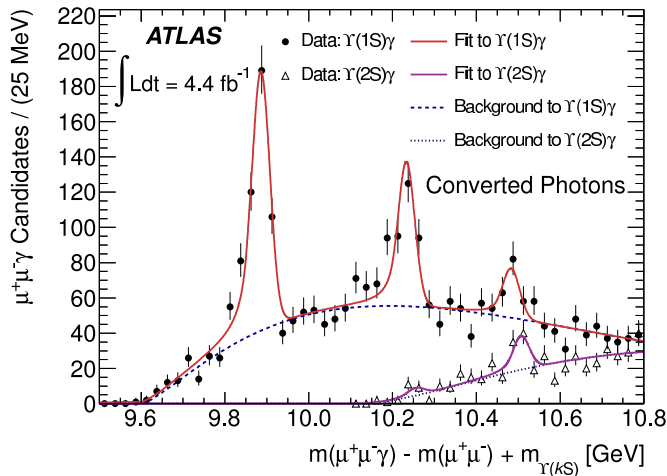
**Figure 3:** From Belle [34]  $e^+e^-$  collision data taken near the peak of the  $Y(10860)$ , the  $h_b(1P)$  event yield vs. the mass recoiling against the  $\pi^+\pi^-\gamma$  (corrected for misreconstructed  $\pi^+\pi^-$ ), where the  $h_b(1P)$  yield is obtained by fitting the mass recoiling against the  $\pi^+\pi^-$  (points with error bars). The fit results (solid histograms) for signal plus background and background alone are superimposed.

samples. Belle fits all subsamples to resonant plus non-resonant amplitudes, allowing for interference (notably, between  $\pi^- Z_b^+$  and  $\pi^+ Z_b^-$ ), and finds consistent pairs of  $Z_b^+$  masses for all bottomonium transitions, and comparable strengths of the two states. A recent angular analysis assigned  $J^P = 1^+$  for both  $Z_b^+$

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states [72], which must also have negative  $G$ -parity. Transitions through  $Z_b^+$  to the  $h_b(nP)$  saturate the observed  $\pi^+\pi^-h_b(nP)$  cross sections. While the two masses of the  $Z_b^+$  states as extracted from Breit-Wigner fits for the various channels are just a few MeV above the  $B^*\bar{B}$  and  $B^*\bar{B}^*$  thresholds, respectively, more refined analyses find pole locations right below the corresponding thresholds either on the physical [118] or the unphysical sheet [119]. Regardless their proximity to the corresponding thresholds, both states predominantly decay into these open flavor channels [74], regardless the small phase space, with branching fractions that exceed 80% and 70%, respectively, at 90% CL. This feature provides strong evidence for their molecular nature—note that the  $Z_b^+$  states cannot be simple mesons because they are charged and have  $b\bar{b}$  content.

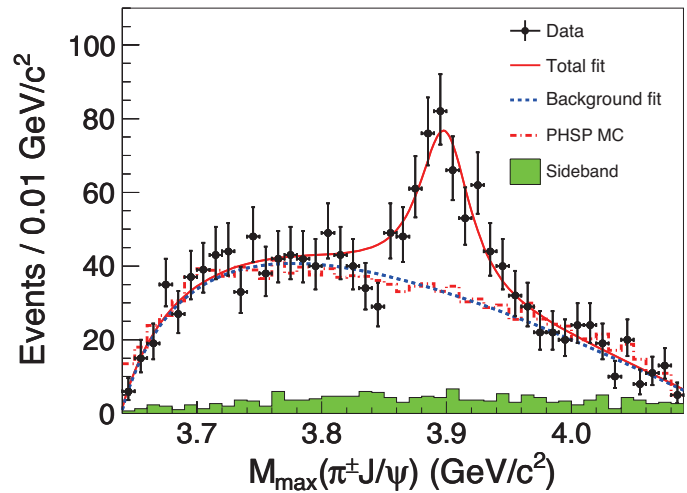


**Figure 4:** From ATLAS [39]  $pp$  collision data (points with error bars) taken at  $\sqrt{s} = 7$  TeV, the effective mass of  $\chi_{bJ}(1P, 2P, 3P) \rightarrow \gamma Y(1S, 2S)$  candidates in which  $Y(1S, 2S) \rightarrow \mu^+\mu^-$  and the photon is reconstructed as an  $e^+e^-$  conversion in the tracking system. Fits (smooth curves) show significant signals for each triplet (merged- $J$ ) on top of a smooth background. From [39] with kind permission, copyright (2012) The American Physical Society.

The third Belle result to follow from these data is the confirmation of the  $\eta_b(1S)$  and measurement of the  $h_b(1P) \rightarrow \gamma\eta_b(1S)$  branching fraction, expected to be several tens of percent. To accomplish this, events with the  $\pi^+\pi^-$  recoil mass in the  $h_b(1P)$  mass window and a radiative photon candidate are selected, and the  $\pi^+\pi^-\gamma$  recoil mass queried for correlation with non-zero  $h_b(1P)$  population in the  $\pi^+\pi^-$  missing mass spectrum, as shown in Fig. 3. A clear peak is observed, corresponding to the  $\eta_b(1S)$ . A fit is performed to extract the  $\eta_b(1S)$  mass, and determine its width and the branching fraction for  $h_b(1P) \rightarrow \gamma\eta_b(1S)$

(the latter of which is  $(49.8 \pm 6.8_{-5.2}^{+10.9})\%$ ) for the first time. The mass determination has comparable uncertainty and a larger central value (by 10 MeV, or  $2.4\sigma$ ) than the average of previous measurements, thereby reducing the new world average hyperfine splitting by nearly 5 MeV. An independent experimental confirmation of the shifted mass recently came from the Belle observation of the  $Y(4S) \rightarrow \eta h_b(1P)$  [33].

The  $\chi_{bJ}(nP)$  states have recently been observed at the LHC by ATLAS [39] and confirmed by D0 [40] for  $n = 1, 2, 3$ , although in each case the three  $J$  states are not distinguished from one another. Events are sought which have both a photon and an  $Y(1S, 2S) \rightarrow \mu^+\mu^-$  candidate which together form a mass in the  $\chi_b$  region. Observation of all three  $J$ -merged peaks is seen with significance in excess of  $6\sigma$  for both unconverted and converted photons. The mass plot for converted photons, which provide better mass resolution, is shown in Fig. 4. This marks the first observation of the  $\chi_{bJ}(3P)$  triplet, quite near the expected mass. A precise confirmation of this result came from LHCb [41].



**Figure 5:**  $J/\psi\pi$  invariant mass distributions from BES-III [61]  $e^+e^-$  collision data taken near the peak of the  $Y(4260)$ . Adapted from [61] with kind permission, copyright (2013) The American Physical Society.

In 2013 at BESIII [61] and shortly after at Belle [62] a charged state called  $Z_c(3900)^+$  was found near the  $D\bar{D}^*$  threshold—the corresponding spectrum from BESIII is shown in Fig. 5. In addition to confirming these findings, Ref. [63] also provided evidence for a neutral partner. A nearby signal was also seen in the  $D\bar{D}^*$  channel [65] whose quantum numbers were fixed to  $1^{+-}$ . The masses extracted from these experiments agree only within  $2\sigma$ . However, since the extraction did not allow for an interference with the background and used Breit-Wigner line shapes, which is not justified near thresholds, there might be some additional systematic uncertainty in the mass values. Therefore in the RPP listings as well as Table 2



both structures appear under the name  $Z_c(3900)^+$ . Analogously,  $Z_c(4020)^+$  (seen in in  $h_c\pi\pi$  [67]) and  $Z_c^+(4025)$  (seen in  $D^*\bar{D}^*$  [69]) are listed as one state,  $Z_c(4020)^+$ . The  $Z_c^+$  states show some remarkable similarities to the  $Z_b^+$  states, e.g. they decay dominantly to the  $D^{(*)}\bar{D}^*$  channels. However, current analyses suggest that the mass of especially the  $Z_c(3900)^+$  might be somewhat above the  $D\bar{D}^*$  threshold. If confirmed, this feature would clearly challenge a possible  $D\bar{D}^*$ -molecular interpretation.

## References

- N. Brambilla *et al.*, CERN-2005-005, (CERN, Geneva, 2005), [arXiv:hep-ph/0412158](#).
- E. Eichten *et al.*, Rev. Mod. Phys. **80**, 1161 (2008) [arXiv:hep-ph/0701208](#).
- S. Eidelman, H. Mahlke-Kruger, and C. Patrignani, in C. Amsler *et al.* (Particle Data Group), Phys. Lett. **B667**, 1029 (2008).
- S. Godfrey and S.L. Olsen, Ann. Rev. Nucl. Part. Sci. **58** 51 (2008), [arXiv:0801.3867 \[hep-ph\]](#).
- T. Barnes and S.L. Olsen, Int. J. Mod. Phys. **A24**, 305 (2009).
- G.V. Pakhlova, P.N. Pakhlov, and S.I. Eidelman, Phys. Usp. **53**, 219 (2010), [Usp. Fiz. Nauk **180**, 225 (2010)].
- N. Brambilla *et al.*, Eur. Phys. J. **C71**, 1534 (2011) [arXiv:1010.5827 \[hep-ph\]](#).
- N. Brambilla *et al.*, Eur. Phys. J. **C 74**, 2981 (2014) [arXiv:1404.3723 \[hep-ph\]](#).
- P. Rubin *et al.* (CLEO Collab.), Phys. Rev. **D72**, 092004 (2005) [arXiv:hep-ex/0508037](#).
- J.L. Rosner *et al.* (CLEO Collab.), Phys. Rev. Lett. **95**, 102003 (2005) [arXiv:hep-ex/0505073](#).
- S. Dobbs *et al.* (CLEO Collab.), Phys. Rev. Lett. **101**, 182003 (2008) [arXiv:0805.4599 \[hep-ex\]](#).
- M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **104**, 132003 (2010) [arXiv:1002.0501 \[hep-ex\]](#).
- M. Andreotti *et al.* (E835 Collab.), Phys. Rev. **D72**, 032001 (2005).
- M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. D **86**, 092009 (2012) [arXiv:1209.4963 \[hep-ex\]](#).
- S.K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **89**, 102001 (2002) [Erratum-ibid. **89**, 129901 (2002)], [arXiv:hep-ex/0206002](#).
- A. Vinokurova *et al.* (Belle Collab.), Phys. Lett. **B706**, 139 (2011) [arXiv:1105.0978 \[hep-ex\]](#).
- P. del Amo Sanchez *et al.* (BaBar Collab.), Phys. Rev. **D84**, 021004 (2011) [arXiv:1103.3971 \[hep-ex\]](#).
- B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **92**, 142002 (2004) [arXiv:hep-ex/0311038](#).
- D.M. Asner *et al.* (CLEO Collab.), Phys. Rev. Lett. **92**, 142001 (2004) [arXiv:hep-ex/0312058](#).
- H. Nakazawa (Belle Collab.), Nucl. Phys. (Proc. Supp.) **184**, 220 (2008).
- B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D72**, 031101 (2005) [arXiv:hep-ex/0506062](#).
- K. Abe *et al.* (Belle Collab.), Phys. Rev. Lett. **98**, 082001 (2007) [arXiv:hep-ex/0507019](#).
- V. Bhardwaj *et al.* (Belle Collaboration), Phys. Rev. Lett. **111**, 032001 (2013), [arXiv:1304.3975](#).
- M. Ablikim *et al.*, Phys. Rev. Lett. **115**, 011803 (2015), [arXiv:1503.08203 \[hep-ex\]](#).
- F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **81**, 2432 (1998) [arXiv:hep-ex/9805034](#).
- T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **100**, 182002 (2008) [arXiv:0712.1506 \[hep-ex\]](#).
- V.M. Abazov *et al.* (D0 Collab.), Phys. Rev. Lett. **101**, 012001 (2008) [arXiv:0802.4258 \[hep-ex\]](#).
- G. Aad *et al.* (ATLAS Collab.), Phys. Rev. Lett. **113**, 212004 (2014), [arXiv:1407.1032 \[hep-ex\]](#).
- B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **101**, 071801 (2008) [Erratum-ibid. **102**, 029901 (2009)], [arXiv:0807.1086 \[hep-ex\]](#).
- G. Bonvicini *et al.* (CLEO Collab.), Phys. Rev. **D81**, 031104 (2010) [arXiv:0909.5474 \[hep-ex\]](#).
- B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **103**, 161801 (2009) [arXiv:0903.1124 \[hep-ex\]](#).
- R. Mizuk *et al.* (Belle Collaboration), Phys. Rev. Lett. **109**, 232002 (2012), [arXiv:1205.6351](#).
- U. Tamponi *et al.* (Belle Collab.), Phys. Rev. Lett. **115**, 142001 (2015), [arXiv:1506.08914 \[hep-ex\]](#).
- I. Adachi *et al.* (Belle Collab.), [arXiv:1110.3934 \[hep-ex\]](#).
- I. Adachi *et al.* (Belle Collab.), Phys. Rev. Lett. **108**, 032001 (2012) [arXiv:1103.3419 \[hep-ex\]](#).
- J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D84**, 091101 (2011) [arXiv:1102.4565 \[hep-ex\]](#).
- G. Bonvicini *et al.* (CLEO Collab.), Phys. Rev. **D70**, 032001 (2004) [arXiv:hep-ex/0404021](#).
- P. del Amo Sanchez *et al.* (BaBar Collab.), Phys. Rev. **D82**, 111102 (2010) [arXiv:1004.0175 \[hep-ex\]](#).
- G. Aad *et al.* (ATLAS Collab.), Phys. Rev. Lett. **108**, 152001 (2012) [arXiv:1112.5154 \[hep-ex\]](#).
- V. M. Abazov (D0 Collab.), Phys. Rev. D **86**, 031103 (2012) [arXiv:1203.6034](#).
- R. Aaij *et al.* (LHCb Collab.), JHEP **1410**, 088 (2014), [arXiv:1409.1408 \[hep-ex\]](#).
- S.K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **91**, 262001 (2003) [arXiv:hep-ex/0309032](#).
- S.-K. Choi *et al.* (Belle Collab.), Phys. Rev. **D84**, 052004R (2011) [arXiv:1107.0163 \[hep-ex\]](#).
- B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D77**, 111101 (2008) [arXiv:0803.2838 \[hep-ex\]](#).
- D.E. Acosta *et al.* (CDF Collab.), Phys. Rev. Lett. **93**, 072001 (2004) [arXiv:hep-ex/0312021](#).
- A. Abulencia *et al.* (CDF Collab.), Phys. Rev. Lett. **98**, 132002 (2007) [arXiv:hep-ex/0612053](#).
- T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **103**, 152001 (2009) [arXiv:0906.5218 \[hep-ex\]](#).
- V.M. Abazov *et al.* (D0 Collab.), Phys. Rev. Lett. **93**, 162002 (2004) [arXiv:hep-ex/0405004](#).
- K. Abe *et al.* (Belle Collab.), [arXiv:hep-ex/0505037](#).
- P. del Amo Sanchez *et al.* (BaBar Collab.), Phys. Rev. **D82**, 011101R (2010) [arXiv:1005.5190 \[hep-ex\]](#).
- G. Gokhroo *et al.* (Belle Collab.), Phys. Rev. Lett. **97**, 162002 (2006) [arXiv:hep-ex/0606055](#).
- T. Aushev *et al.* (Belle Collab.), Phys. Rev. **D81**, 031103R (2010), [arXiv:0810.0358 \[hep-ex\]](#).

# Meson Particle Listings

## Heavy Quarkonium Spectroscopy

53. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D77**, 011102 (2008) [arXiv:0708.1565 \[hep-ex\]](#).
54. V. Bhardwaj *et al.* (Belle Collab.), Phys. Rev. Lett. **107**, 091803 (2011) [arXiv:1105.0177 \[hep-ex\]](#).
55. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D74**, 071101 (2006) [arXiv:hep-ex/0607050](#).
56. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **102**, 132001 (2009) [arXiv:0809.0042 \[hep-ex\]](#).
57. R. Aaij *et al.* (LHCb Collab.), Nucl. Phys. B **886**, 665 (2014) [arXiv:1404.0275 \[hep-ex\]](#).
58. R. Aaij *et al.* (LHCb Collab.), Eur. Phys. J. **C72**, 1972 (2012) [arXiv:1112.5310 \[hep-ex\]](#).
59. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **110**, 222001 (2013) [arXiv:1302.6269 \[hep-ex\]](#).
60. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. D **92**, 011102 (2015) [arXiv:1504.06339 \[hep-ex\]](#).
61. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **110**, 252001 (2013) [arXiv:1303.5949](#).
62. Z. Q. Liu *et al.* (Belle Collab.), Phys. Rev. Lett. **110**, 252002 (2013) [arXiv:1304.0121](#).
63. T. Xiao, S. Dobbs, A. Tomaradze and K. K. Seth, Phys. Lett. **B727**, 366 (2013) [arXiv:1304.3036](#).
64. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **115**, 112003 (2015) [arXiv:1506.06018 \[hep-ex\]](#).
65. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **112**, 022001 (2014) [arXiv:1310.1163](#).
66. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **115**, 222002 (2015) [arXiv:1509.05620 \[hep-ex\]](#).
67. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **111**, 242001 (2013) [arXiv:1309.1896](#).
68. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **113**, 212002 (2014) [arXiv:1409.6577 \[hep-ex\]](#).
69. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **112**, 132001 (2014), [arXiv:1308.2760](#).
70. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **115**, 182002 (2015) [arXiv:1507.02404 \[hep-ex\]](#).
71. A. Bondar *et al.* (Belle Collab.), Phys. Rev. Lett. **108**, 122001 (2012) [arXiv:1110.2251 \[hep-ex\]](#).
72. A. Garmash *et al.* (Belle Collab.), Phys. Rev. D **91**, 072003 (2015) [arXiv:1403.0992](#).
73. P. Krokovny *et al.* (Belle Collab.), Phys. Rev. D **88**, 052016 (2013), [arXiv:1308.2646](#).
74. I. Adachi *et al.* (Belle Collab.) (2012), [arXiv:1209.6450](#).
75. S.-K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **94**, 182002 (2005) [arXiv:hep-ex/0408126](#).
76. S. Uehara *et al.* (Belle Collab.), Phys. Rev. Lett. **104**, 092001 (2010) [arXiv:0912.4451 \[hep-ex\]](#).
77. J. P. Lees *et al.* (BaBar Collab.), Phys. Rev. D **86**, 072002 (2012) [arXiv:1207.2651 \[hep-ex\]](#).
78. S. Uehara *et al.* (Belle Collab.), Phys. Rev. Lett. **96**, 082003 (2006) [arXiv:hep-ex/0512035](#).
79. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. **D81**, 092003 (2010) [arXiv:1002.0281 \[hep-ex\]](#).
80. P. Pakhlov *et al.* (Belle Collab.), Phys. Rev. Lett. **100**, 202001 (2008) [arXiv:0708.3812 \[hep-ex\]](#).
81. C.Z. Yuan *et al.* (Belle Collab.), Phys. Rev. Lett. **99**, 182004 (2007) [arXiv:0707.2541 \[hep-ex\]](#).
82. R. Mizuk *et al.* (Belle Collab.), Phys. Rev. **D78**, 072004 (2008) [arXiv:0806.4098 \[hep-ex\]](#).
83. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. **D85**, 052003 (2011) [arXiv:1111.5919 \[hep-ex\]](#).
84. T. Aaltonen *et al.* (CDF Collab.), Phys. Rev. Lett. **102**, 242002 (2009) [arXiv:0903.2229 \[hep-ex\]](#).
85. T. Aaltonen *et al.* (CDF Collab.), [arXiv:1101.6058 \[hep-ex\]](#).
86. V. Abazov *et al.* (D0 Collab.), Phys. Rev. D **89**, 012004 (2014), [arXiv:1309.6580](#).
87. S. Chatrchyan *et al.* (CMS Collab.), Phys. Lett. B **734**, 261 (2014) [arXiv:1309.6920](#).
88. J. Brodzicka (Belle Collab.), Conf. Proc. C0908171, 299 (2009).
89. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. D **85**, 091103 (2012), [arXiv:1202.5087](#).
90. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. D **91**, 012003 (2015), [arXiv:1407.7244](#).
91. C.P. Shen *et al.* (Belle Collab.), Phys. Rev. Lett. **104**, 112004 (2010) [arXiv:0912.2383 \[hep-ex\]](#).
92. K. Chilikin *et al.* (Belle Collab.), Phys. Rev. D **90**, 112009 (2014) [arXiv:1408.6457 \[hep-ex\]](#).
93. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **95**, 142001 (2005) [arXiv:hep-ex/0506081](#).
94. B. Aubert *et al.* (BaBar Collab.), [arXiv:0808.1543v2 \[hep-ex\]](#).
95. Q. He *et al.* (CLEO Collab.), Phys. Rev. **D74**, 091104 (2006) [arXiv:hep-ex/0611021](#).
96. T.E. Coan *et al.* (CLEO Collab.), Phys. Rev. Lett. **96**, 162003 (2006) [arXiv:hep-ex/0602034](#).
97. J. Lees *et al.* (BaBar Collab.), Phys. Rev. D **86**, 051102 (2012), [arXiv:1204.2158](#).
98. M. Ablikim *et al.* (BESIII Collab.), Phys. Rev. Lett. **112**, 092001 (2014), [arXiv:1310.4101](#).
99. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. Lett. **98**, 212001 (2007) [arXiv:hep-ex/0610057](#).
100. J.P. Lees *et al.* (BaBar Collab.), Phys. Rev. D **89**, 111103 (2014) [arXiv:1211.6271 \[hep-ex\]](#).
101. X.L. Wang *et al.* (Belle Collab.), Phys. Rev. Lett. **99**, 142002 (2007) [arXiv:0707.3699 \[hep-ex\]](#).
102. X.L. Wang *et al.* (Belle Collab.), Phys. Rev. D **91**, 112007 (2015) [arXiv:1410.7641 \[hep-ex\]](#).
103. S.K. Choi *et al.* (Belle Collab.), Phys. Rev. Lett. **100**, 142001 (2008) [arXiv:0708.1790 \[hep-ex\]](#).
104. R. Mizuk *et al.* (Belle Collab.), Phys. Rev. **D80**, 031104 (2009) [arXiv:0905.2869 \[hep-ex\]](#).
105. K. Chilikin *et al.* (Belle Collab.), Phys. Rev. D **88**, 074026 (2013), [arXiv:1306.4894](#).
106. B. Aubert *et al.* (BaBar Collab.), Phys. Rev. D **79**, 112001 (2009), [arXiv:0811.0564](#).
107. R. Aaij *et al.* (LHCb Collab.), Phys. Rev. Lett. **112** no.22, 222002 (2014) [arXiv:1404.1903 \[hep-ex\]](#).
108. G. Pakhlova *et al.* (Belle Collab.), Phys. Rev. Lett. **101**, 172001 (2008) [arXiv:0807.4458 \[hep-ex\]](#).
109. K. Olive *et al.* (Particle Data Group), Chin. Phys. C **38**, 090001 (2014).
110. K.F. Chen *et al.* (Belle Collab.), Phys. Rev. Lett. **100**, 112001 (2008) [arXiv:0710.2577 \[hep-ex\]](#).
111. D. Santel *et al.* (Belle Collab.), Phys. Rev. D **93**, 011101 (2016) [arXiv:1501.01137 \[hep-ex\]](#).

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| <p>112. P. Krokovny (Belle Collab.), talk given at Les Rencontres de Physique de la Vallee d'Aoste, La Thuile, Aosta Valley, Italy, 2012.</p> <p>113. A. Abdesselam <i>et al.</i> (Belle Collab.), <a href="#">arXiv:1508.06562</a>.</p> <p>114. F.-K. Guo and U.-G. Meissner, Phys. Rev. D <b>86</b>, 091501 (2012) <a href="#">arXiv:1208.1134</a>.</p> <p>115. S.L. Olsen, Phys. Rev. D <b>91</b>, 057501 (2015) <a href="#">arXiv:1410.6534</a>.</p> | <p>116. Z. Y. Zhou, Z. Xiao and H. Q. Zhou, Phys. Rev. Lett. <b>115</b> (2015) no.2, 022001 <a href="#">arXiv:1501.00879 [hep-ph]</a>.</p> <p>117. T.K. Pedlar <i>et al.</i> (CLEO Collab.), Phys. Rev. Lett. <b>107</b>, 041803 (2011) <a href="#">arXiv:1104.2025 [hep-ex]</a>.</p> <p>118. M. Cleven <i>et al.</i>, Eur. Phys. J. A <b>47</b>, 120 (2011) <a href="#">arXiv:1107.0254 [hep-ph]</a>.</p> <p>119. F.-K. Guo <i>et al.</i>, <a href="#">arXiv:1602.00940 [hep-ph]</a>.</p> |
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## $c\bar{c}$ MESONS

[illegible]

The level scheme of the  $c\bar{c}$  states showing experimentally established states with solid lines. Singlet states are called  $\eta_c$  and  $h_c$ , triplet states  $\psi$  and  $\chi_{cJ}$ , and unassigned charmonium-like states  $X$ . In parentheses it is sufficient to give the radial quantum number and the orbital angular momentum to specify the states with all their quantum numbers. Only observed hadronic transitions are shown; the single photon transitions  $\psi(nS) \rightarrow \gamma\eta_c(mP)$ ,  $\psi(nS) \rightarrow \gamma\chi_{cJ}(mP)$ , and  $\chi_{cJ}(1P) \rightarrow \gamma J/\psi$  are omitted for clarity.

 $\eta_c(1S)$ 

$$1^G(J^{PC}) = 0^+(0^--)$$

### $\eta_c(1S)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2983.4 ± 0.5</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.2.		
2982.2 ± 1.5 ± 0.1	2.0k	1 AAJ	15BI LHCB	$p\bar{p} \rightarrow \eta_c(1S)X$
2983.5 ± 1.4 ± 1.6		2 ANASHIN	14 KEDR	$J/\psi \rightarrow \gamma\eta_c$
2979.8 ± 0.8 ± 3.5	4.5k	3,4 LEES	14E BABR	$\gamma\gamma \rightarrow K^+K^-\pi^0$
2984.1 ± 1.1 ± 2.1	900	3,4,5 LEES	14E BABR	$\gamma\gamma \rightarrow K^+K^-\eta$
2984.3 ± 0.6 ± 0.6		6,7 ABLIKIM	12F BES3	$\psi(2S) \rightarrow \gamma\eta_c$
2984.49 ± 1.16 ± 0.52	832	3 ABLIKIM	12N BES3	$\psi(2S) \rightarrow \pi^0\gamma \text{ hadrons}$
2982.7 ± 1.8 ± 2.2	486	ZHANG	12A BELL	$e^+e^- \rightarrow$ $e^+e^-\eta/\pi^+\pi^-$
2984.5 ± 0.8 ± 3.1	11k	DEL-AMO-SA...	11M BABR	$\gamma\gamma \rightarrow$ $K^+K^-\pi^+\pi^-\pi^0$
2985.4 ± 1.5 ± 0.5	920	7 VINOKUROVA	11 BELL	$B^{\pm} \rightarrow$ $K^{\pm}(K_S^0 K^{\pm}\pi^{\mp})$
2982.2 ± 0.4 ± 1.6	14k	8 LEES	10 BABR	$10.6 e^+e^- \rightarrow$ $K_S^0 K^{\pm}\pi^{\mp}$
2985.8 ± 1.5 ± 3.1	0.9k	AUBERT	08AB BABR	$B \rightarrow \eta_c(1S) K(*) \rightarrow$ $K\bar{K}\pi K(*)$
2986.1 ± 1.0 ± 2.5	7.5k	UEHARA	08 BELL	$\gamma\gamma \rightarrow \eta_c \rightarrow \text{hadrons}$
2970 ± 5 ± 6	501	9 ABE	07 BELL	$e^+e^- \rightarrow J/\psi(c\bar{c})$
2971 ± 3 ± 2	195	WU	06 BELL	$B^+ \rightarrow p\bar{p}K^+$
2974 ± 7 ± 2	20	WU	06 BELL	$B^+ \rightarrow \Lambda\bar{\Lambda}K^+$
2981.8 ± 1.3 ± 1.5	592	ASNER	04 CLEO	$\gamma\gamma \rightarrow \eta_c \rightarrow$ $K_S^0 K^{\pm}\pi^{\mp}$
2984.1 ± 2.1 ± 1.0	190	10 AMBROGIANI	03 E835	$\bar{p}p \rightarrow \eta_c \rightarrow \gamma\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2982.5 ± 0.4 ± 1.4	12k	11 DEL-AMO-SA...	11M BABR	$\gamma\gamma \rightarrow K_S^0 K^{\pm}\pi^{\mp}$
2982.2 ± 0.6		12 MITCHELL	09 CLEO	$e^+e^- \rightarrow \gamma X$
2982 ± 5	270	13 AUBERT	06E BABR	$B^{\pm} \rightarrow K^{\pm}X_{c\bar{c}}$
2982.5 ± 1.1 ± 0.9	2.5k	14 AUBERT	04D BABR	$\gamma\gamma \rightarrow \eta_c(1S) \rightarrow$ $K\bar{K}\pi$

2977.5 $\pm$ 1.0 $\pm$ 1.2		<sup>12,15</sup> BAI	03	BES	$J/\psi \rightarrow \gamma \eta_C$
2979.6 $\pm$ 2.3 $\pm$ 1.6	180	<sup>16</sup> FANG	03	BELL	$B \rightarrow \eta_C K$
2976.3 $\pm$ 2.3 $\pm$ 1.2		<sup>12,17</sup> BAI	00F	BES	$J/\psi, \psi(2S) \rightarrow \gamma \eta_C$
2976.6 $\pm$ 2.9 $\pm$ 1.3	140	<sup>12,18</sup> BAI	00F	BES	$J/\psi \rightarrow \gamma \eta_C$
2980.4 $\pm$ 2.3 $\pm$ 0.6		<sup>19</sup> BRANDENB...	00B	CLE2	$\gamma \gamma \rightarrow \eta_C \rightarrow K^\pm K_S^0 \pi^\mp$
2975.8 $\pm$ 3.9 $\pm$ 1.2		<sup>18</sup> BAI	99B	BES	Sup. by BAI 00F
2999 $\pm$ 8	25	ABREU	98a	DLPH	$e^+ e^- \rightarrow e^+ e^- + \text{hadrons}$
2988.3 $\pm$ $\begin{smallmatrix} 3.3 \\ -3.1 \end{smallmatrix}$		ARMSTRONG	95F	E760	$\bar{p} p \rightarrow \gamma \gamma$
2974.4 $\pm$ 1.9		<sup>12,20</sup> BISELLO	91	DM2	$J/\psi \rightarrow \eta_C \gamma$
2969 $\pm$ 4 $\pm$ 4	80	<sup>12</sup> BAI	90B	MRK3	$J/\psi \rightarrow \gamma K^+ K^- K^+ K^-$
2956 $\pm$ 12 $\pm$ 12		<sup>12</sup> BAI	90B	MRK3	$J/\psi \rightarrow \gamma K^+ K^- K_S^0 K_L^0$
2982.6 $\pm$ $\begin{smallmatrix} 2.7 \\ -2.3 \end{smallmatrix}$	12	BAGLIN	87B	SPEC	$\bar{p} p \rightarrow \gamma \gamma$
2980.2 $\pm$ 1.6		<sup>12,20</sup> BALTRUSAIT...	86	MRK3	$J/\psi \rightarrow \eta_C \gamma$
2984 $\pm$ 2.3 $\pm$ 4.0		<sup>12</sup> GAISER	86	CBAL	$J/\psi \rightarrow \gamma X, \psi(2S) \rightarrow \gamma X$
2976 $\pm$ 8		<sup>12,21</sup> BALTRUSAIT...	84	MRK3	$J/\psi \rightarrow 2 \phi \gamma$
2982 $\pm$ 8	18	<sup>22</sup> HIMEL	80B	MRK2	$e^+ e^-$
2980 $\pm$ 9		<sup>22</sup> PARTRIDGE	80B	CBAL	$e^+ e^-$

<sup>1</sup>AAJ15BI reports  $m_{J/\psi} - m_{\eta_c(1S)} = 114.7 \pm 1.5 \pm 0.1$  MeV from a sample of  $\eta_c(1S)$  and  $J/\psi$  produced in  $b$ -hadron decays. We have used current value of  $m_{J/\psi} = 3096.900 \pm 0.006$  MeV to arrive at the quoted  $m_{\eta_c(1S)}$  result.

<sup>2</sup>Taking into account an asymmetric photon lineshape.

<sup>3</sup>With floating width.

<sup>4</sup> Ignoring possible interference with the non-resonant  $0^-$  amplitude.

<sup>5</sup> Using both,  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decays.

<sup>6</sup>From a simultaneous fit to six decay modes of the  $\eta_C$ .

<sup>7</sup>Accounts for interference with non-resonant continuum.

<sup>8</sup> Taking into account interference with the

<sup>9</sup>From a fit of the  $J/\psi$  recoil mass spectrum. Supersedes ABE,k

<sup>10</sup> Using mass of  $\psi(2S) = 3686.00$  MeV.  
<sup>11</sup> Not independent from the measurements reported by LEES <sup>10</sup>.  
<sup>12</sup> NOT INDEPENDENT FROM THE MEASUREMENTS REPORTED BY LEES <sup>10</sup>.

<sup>12</sup> MITCHELL 09 observes a significant asymmetry in the lineshapes of  $\psi(2S) \rightarrow \psi(1S)\gamma$  and  $\psi(3S) \rightarrow \psi(1S)\gamma$  transitions. If indeed, this asymmetry could lead to signif-

and  $J/\psi \rightarrow \gamma \eta_C$  transition  
whenever the mass and width

<sup>13</sup>From the fit of the kaon momentum spectrum. Systematic errors not evaluated.

14 Superseded by LEFS 1.0

<sup>15</sup>From a simultaneous fit of five decay modes of the  $\eta_c$ .

<sup>16</sup> Superseded by VINOKUROVA 11.

<sup>17</sup>Weighted average of the  $\psi(2S)$  and  $J/\psi(1S)$  samples. Using an  $\eta_c$  width of 13.2 MeV.

<sup>18</sup> Average of several decay modes. Using an  $\eta_c$  width of 13.2 MeV.

<sup>19</sup> Superseded by ASNER 04.

<sup>20</sup> Average of several decay modes.

$$^{21}\eta_c \rightarrow \phi\phi.$$

<sup>22</sup> Mass adjusted by us to correspond to  $J/\psi(1S)$  mass = 3097 MeV.

### $\eta_c(1S)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>31.9± 0.8 OUR FIT</b>				
<b>31.9± 1.0 OUR AVERAGE</b>				Error includes scale factor of 1.2.
27.2± 3.1 <sup>+5.4</sup> <sub>-2.6</sub>		<sup>1</sup> ANASHIN	14 KEDR	$J/\psi \rightarrow \gamma \eta_C$
25.2± 2.6± 2.4	4.5k	<sup>2,3</sup> LEES	14E BABR	$\gamma\gamma \rightarrow K^+ K^- \pi^0$
34.8± 3.1± 4.0	900	<sup>2,3,4</sup> LEES	14E BABR	$\gamma\gamma \rightarrow K^+ K^- \eta$
32.0± 1.2± 1.0		<sup>5,6</sup> ABLIKIM	12F BES3	$\psi(2S) \rightarrow \gamma \eta_C$
36.4± 3.2± 1.7	832	<sup>2</sup> ABLIKIM	12N BES3	$\psi(2S) \rightarrow \pi^0 \gamma \text{ hadrons}$
37.8 <sup>+5.8</sup> <sub>-5.3</sub> ± 3.1	486	ZHANG	12A BELL	$e^+ e^- \rightarrow e^+ e^- \eta' \pi^+ \pi^-$
36.2± 2.8± 3.0	11k	DEL-AMO-SA...	11M BABR	$\gamma\gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$
35.1± 3.1 <sup>+1.0</sup> <sub>-1.6</sub>	920	<sup>6</sup> VINOKUROVA	11 BELL	$B^\pm \rightarrow K^\pm (K_S^0 K^\pm \pi^\mp)$
31.7± 1.2± 0.8	14k	<sup>7</sup> LEES	10 BABR	$10.6 e^+ e^- \rightarrow K_S^0 K^\pm \pi^\mp$
36.3 <sup>+3.7</sup> <sub>-3.6</sub> ± 4.4	0.9k	AUBERT	08AB BABR	$B \rightarrow \eta_C(1S) K^*(*) \rightarrow K \bar{K} \pi K^*(*)$
28.1± 3.2± 2.2	7.5k	UEHARA	08 BELL	$\gamma\gamma \rightarrow \eta_C \rightarrow \text{hadrons}$
48 <sup>+8</sup> <sub>-7</sub> ± 5	195	WU	06 BELL	$B^+ \rightarrow p \bar{p} K^+$
40 ± 19 ± 5	20	WU	06 BELL	$B^+ \rightarrow \Lambda \bar{\Lambda} K^+$
24.8± 3.4± 3.5	592	ASNER	04 CLEO	$\gamma\gamma \rightarrow \eta_C \rightarrow K_S^0 K^\pm \pi^\mp$
20.4 <sup>+7.7</sup> <sub>-6.7</sub> ± 2.0	190	AMBROGIANI	03 E835	$\bar{p} p \rightarrow \eta_C \rightarrow \gamma\gamma$
23.9 <sup>+12.6</sup> <sub>-7.1</sub>		ARMSTRONG	95F E760	$\bar{p} p \rightarrow \gamma\gamma$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
32.1± 1.1± 1.3	12k	<sup>8</sup> DEL-AMO-SA...	11M BABR	$\gamma\gamma \rightarrow K_S^0 K^\pm \pi^\mp$
34.3± 2.3± 0.9	2.5k	<sup>9</sup> AUBERT	04D BABR	$\gamma\gamma \rightarrow \eta_C(1S) \rightarrow K \bar{K} \pi$
17.0± 3.7± 7.4		<sup>10</sup> BAI	03 BES	$J/\psi \rightarrow \gamma \eta_C$
29 ± 8 ± 6	180	<sup>11</sup> FANG	03 BELL	$B \rightarrow \eta_C K$
11.0± 8.1± 4.1		<sup>12</sup> BAI	00F BES	$J/\psi \rightarrow \gamma \eta_C \text{ and } \psi(2S) \rightarrow \gamma \eta_C$
27.0± 5.8± 1.4		<sup>13</sup> BRANDENB...	00B CLE2	$\gamma\gamma \rightarrow \eta_C \rightarrow K^\pm K_S^0 \pi^\mp$
7.0 <sup>+7.5</sup> <sub>-7.0</sub>	12	BAGLIN	87B SPEC	$\bar{p} p \rightarrow \gamma\gamma$
10.1 <sup>+33.0</sup> <sub>-8.2</sub>	23	<sup>14</sup> BALTRUSAIT...	.86 MRK3	$J/\psi \rightarrow \gamma p \bar{p}$
11.5± 4.5		GAISER	86 CBAL	$J/\psi \rightarrow \gamma X, \psi(2S) \rightarrow \gamma X$
< 40 90% CL	18	HIMEL	80B MRK2	$e^+ e^-$
< 20 90% CL		PARTRIDGE	80B CBAL	$e^+ e^-$

<sup>1</sup> Taking into account an asymmetric photon lineshape.<sup>2</sup>With floating mass.

<sup>3</sup> Ignoring possible interference with the non-resonant  $0^-$  amplitude.

<sup>4</sup> Using both,  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays.

<sup>5</sup> From a simultaneous fit to six decay modes of the  $\eta_c$ .

<sup>6</sup> Accounts for interference with non-resonant continuum.

<sup>7</sup> Taking into account interference with the non-resonant  $J^P = 0^-$  amplitude.

<sup>8</sup> Not independent from the measurements reported by LEES 10.

<sup>9</sup> Superseded by LEES 10.

<sup>10</sup>From a simultaneous fit of five decay modes of the  $\eta_c$ .

<sup>11</sup> Superseded by VINOKUROVA 11.

<sup>12</sup>From a fit to the 4-prong invariant mass in  $\psi(2S) \rightarrow \gamma\eta_c$  and  $J/\psi(1S) \rightarrow \gamma\eta_c$  decays.

<sup>13</sup> Superseded by ASNER 04.

<sup>14</sup> Positive and negative errors correspond to 90% confidence level.

## $\eta_c(1S)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
<b>Decays involving hadronic resonances</b>		
$\Gamma_1$ $\eta'(958)\pi\pi$	( 4.1 $\pm$ 1.7 ) %	
$\Gamma_2$ $\rho\rho$	( 1.8 $\pm$ 0.5 ) %	
$\Gamma_3$ $K^*(892)^0 K^- \pi^+ + \text{c.c.}$	( 2.0 $\pm$ 0.7 ) %	
$\Gamma_4$ $K^*(892) \bar{K}^*(892)$	( 7.0 $\pm$ 1.3 ) $\times 10^{-3}$	
$\Gamma_5$ $K^*(892)^0 \bar{K}^*(892)^0 \pi^+ \pi^-$	( 1.1 $\pm$ 0.5 ) %	
$\Gamma_6$ $\phi K^+ K^-$	( 2.9 $\pm$ 1.4 ) $\times 10^{-3}$	
$\Gamma_7$ $\phi\phi$	( 1.75 $\pm$ 0.20 ) $\times 10^{-3}$	
$\Gamma_8$ $\phi 2(\pi^+ \pi^-)$	< 4 $\times 10^{-3}$	90%
$\Gamma_9$ $a_0(980)\pi$	< 2 %	90%
$\Gamma_{10}$ $a_2(1320)\pi$	< 2 %	90%
$\Gamma_{11}$ $K^*(892) \bar{K} + \text{c.c.}$	< 1.28 %	90%
$\Gamma_{12}$ $f_2(1270)\eta$	< 1.1 %	90%

$\Gamma_{13}$	$\omega\omega$	$< 3.1$	$\times 10^{-3}$	90%
$\Gamma_{14}$	$\omega\phi$	$< 1.7$	$\times 10^{-3}$	90%
$\Gamma_{15}$	$f_2(1270)f_2(1270)$	$(9.8 \pm 2.5)$	$\times 10^{-3}$	
$\Gamma_{16}$	$f_2(1270)f_2'(1525)$	$(9.7 \pm 3.2)$	$\times 10^{-3}$	
$\Gamma_{17}$	$f_0(980)\eta$	seen		
$\Gamma_{18}$	$f_0(1500)\eta$	seen		
$\Gamma_{19}$	$f_0(2200)\eta$	seen		
$\Gamma_{20}$	$a_0(980)\pi$	seen		
$\Gamma_{21}$	$a_0(1320)\pi$	seen		
$\Gamma_{22}$	$a_0(1450)\pi$	seen		
$\Gamma_{23}$	$a_0(1950)\pi$	seen		
$\Gamma_{24}$	$a_2(1950)\pi$	not seen		
$\Gamma_{25}$	$K_0^*(1430)\overline{K}$	seen		
$\Gamma_{26}$	$K_2^*(1430)\overline{K}$	seen		
$\Gamma_{27}$	$K_2^*(1950)\overline{K}$	seen		

### Decays into stable hadrons

$\Gamma_{28}$	$K \bar{K} \pi$	$(7.3 \pm 0.5) \%$
$\Gamma_{29}$	$K \bar{K} \eta$	$(1.35 \pm 0.16) \%$
$\Gamma_{30}$	$\eta \pi^+ \pi^-$	$(1.7 \pm 0.5) \%$
$\Gamma_{31}$	$\eta 2(\pi^+ \pi^-)$	$(4.4 \pm 1.3) \%$
$\Gamma_{32}$	$K^+ K^- \pi^+ \pi^-$	$(6.9 \pm 1.1) \times 10^{-3}$
$\Gamma_{33}$	$K^+ K^- \pi^+ \pi^- \pi^0$	$(3.5 \pm 0.6) \%$
$\Gamma_{34}$	$K^0 K^- \pi^+ \pi^- \pi^+ + \text{c.c.}$	$(5.6 \pm 1.5) \%$
$\Gamma_{35}$	$K^+ K^- 2(\pi^+ \pi^-)$	$(7.5 \pm 2.4) \times 10^{-3}$
$\Gamma_{36}$	$2(K^+ K^-)$	$(1.46 \pm 0.30) \times 10^{-3}$
$\Gamma_{37}$	$\pi^+ \pi^- \pi^0 \pi^0$	$(4.7 \pm 1.0) \%$
$\Gamma_{38}$	$2(\pi^+ \pi^-)$	$(9.7 \pm 1.2) \times 10^{-3}$
$\Gamma_{39}$	$2(\pi^+ \pi^- \pi^0)$	$(17.4 \pm 3.3) \%$
$\Gamma_{40}$	$3(\pi^+ \pi^-)$	$(1.8 \pm 0.4) \%$
$\Gamma_{41}$	$\rho \bar{\rho}$	$(1.50 \pm 0.16) \times 10^{-3}$
$\Gamma_{42}$	$\rho \bar{\rho} \pi^0$	$(3.6 \pm 1.3) \times 10^{-3}$
$\Gamma_{43}$	$\Lambda \bar{\Lambda}$	$(1.09 \pm 0.24) \times 10^{-3}$
$\Gamma_{44}$	$\Sigma^+ \bar{\Sigma}^-$	$(2.1 \pm 0.6) \times 10^{-3}$
$\Gamma_{45}$	$\Xi^- \bar{\Xi}^+$	$(8.9 \pm 2.7) \times 10^{-4}$
$\Gamma_{46}$	$\pi^+ \pi^- \rho \bar{\rho}$	$(5.3 \pm 1.8) \times 10^{-3}$

### Radiative decays

$$\Gamma_{47} \quad \gamma\gamma \quad (1.59 \pm 0.13) \times 10^{-4}$$

Charge conjugation ( $C$ ), Parity ( $P$ ),  
Lepton family number ( $LF$ ) violating modes

$\Gamma_{48}$	$\pi^+ \pi^-$	$P, CP$	$< 1.1$	$\times 10^{-4}$	90%
$\Gamma_{49}$	$\pi^0 \pi^0$	$P, CP$	$< 4$	$\times 10^{-5}$	90%
$\Gamma_{50}$	$K^+ K^-$	$P, CP$	$< 6$	$\times 10^{-4}$	90%
$\Gamma_{51}$	$K_S^0 K_L^0$	$P, CP$	$< 3.1$	$\times 10^{-4}$	90%

### CONSTRAINED FIT INFORMATION

An overall fit to the total width, 8 combinations of partial widths obtained from integrated cross section, and 19 branching ratios uses 85 measurements and one constraint to determine 13 parameters. The overall fit has a  $\chi^2 = 118.3$  for 73 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_7$	18									
$x_{15}$	3	6								
$x_{28}$	22	41	7							
$x_{29}$	12	22	4	54						
$x_{32}$	11	21	4	25	13					
$x_{36}$	9	16	3	25	14	10				
$x_{38}$	14	25	5	30	16	16	12			
$x_{41}$	14	26	5	36	19	16	13	20		
$x_{43}$	3	6	1	9	5	4	3	5	25	
$x_{47}$	-29	-54	-10	-66	-35	-34	-27	-41	-46	-11
$\Gamma$	-2	-3	-1	-4	-2	-2	-1	-2	7	2
	$x_4$	$x_7$	$x_{15}$	$x_{28}$	$x_{29}$	$x_{32}$	$x_{36}$	$x_{38}$	$x_{41}$	$x_{43}$
$\Gamma$	$-28$									
	$x_{47}$									

Mode	Rate (MeV)	
$\Gamma_4$ $K^*(892) \bar{K}^*(892)$	0.22 $\pm$ 0.04	
$\Gamma_7$ $\phi\phi$	0.056 $\pm$ 0.007	
$\Gamma_{15}$ $f_2(1270) f_2(1270)$	0.31 $\pm$ 0.08	
$\Gamma_{28}$ $K \bar{K} \pi$	2.31 $\pm$ 0.16	
$\Gamma_{29}$ $K \bar{K} \eta$	0.43 $\pm$ 0.05	
$\Gamma_{32}$ $K^+ K^- \pi^+ \pi^-$	0.219 $\pm$ 0.034	
$\Gamma_{36}$ $2(K^+ K^-)$	0.046 $\pm$ 0.010	
$\Gamma_{38}$ $2(\pi^+ \pi^-)$	0.31 $\pm$ 0.04	
$\Gamma_{41}$ $\rho \bar{\rho}$	0.048 $\pm$ 0.005	
$\Gamma_{43}$ $\Lambda \bar{\Lambda}$	0.034 $\pm$ 0.008	
$\Gamma_{47}$ $\gamma\gamma$	0.0051 $\pm$ 0.0004	

$\Gamma(\phi\phi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.9 <math>\pm</math> 0.8 OUR FIT</b>				
<b>7.75 <math>\pm</math> 0.66 <math>\pm</math> 0.62</b>	386 $\pm$ 31	<sup>1</sup> LIU	12b BELL	$\gamma\gamma \rightarrow 2(K^+ K^-)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.8 $\pm$ 1.2 $\pm$ 1.3	132 $\pm$ 23	UEHARA	08 BELL	$\gamma\gamma \rightarrow 2(K^+ K^-)$
<sup>1</sup> Supersedes UEHARA 08. Using $B(\phi \rightarrow K^+ K^-) = (48.9 \pm 0.5)\%$ .				

$\Gamma(\omega\omega) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.67 <math>\pm</math> 2.86 <math>\pm</math> 0.96</b>	85 $\pm$ 29	<sup>1</sup> LIU	12b BELL	$\gamma\gamma \rightarrow 2(\pi^+ \pi^- \pi^0)$
<sup>1</sup> Using $B(\omega \rightarrow \pi^+ \pi^- \pi^0) = (89.2 \pm 0.7)\%$ .				

$\Gamma(\omega\phi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$

$\Gamma_{7\Gamma 47}/\Gamma$

$\Gamma_{13\Gamma 47}/\Gamma$

$\Gamma_{14\Gamma 47}/\Gamma$

$\Gamma(\gamma\gamma)$	VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.1<math>\pm</math> 0.4 OUR FIT</b>					
• • •	We do not use the following data for averages, fits, limits, etc. • • •				
5.8 $\pm$ 1.1	486	<sup>1</sup> ZHANG	12A BELL	$e^+e^- \rightarrow e^+e^-\eta'/\pi^+\pi^-$	
5.2 $\pm$ 1.2	273 $\pm$ 43	<sup>2,3</sup> AUBERT	06E BABR	$B^\pm \rightarrow K^\pm X_C \bar{C}$	
5.5 $\pm$ 1.2 $\pm$ 1.8	157 $\pm$ 33	<sup>4</sup> KUO	05 BELL	$\gamma\gamma \rightarrow p\bar{p}$	
7.4 $\pm$ 0.4 $\pm$ 2.3		<sup>5</sup> ASNER	04 CLEO	$\gamma\gamma \rightarrow \eta_C \rightarrow K_S^0 K^\pm \pi^\mp$	
13.9 $\pm$ 2.0 $\pm$ 3.0	41	<sup>6</sup> ABDALLAH	03J DLPH	$\gamma\gamma \rightarrow \eta_C$	
3.8 $\pm$ 1.1 $\pm$ 1.9	190	<sup>7</sup> AMBROGIANI	03 E835	$\bar{p}p \rightarrow \eta_C \rightarrow \gamma\gamma$	
7.6 $\pm$ 0.8 $\pm$ 2.3		<sup>5,8</sup> BRANDENB...	00B CLE2	$\gamma\gamma \rightarrow \eta_C \rightarrow K^\pm K_S^0 \pi^\mp$	
6.9 $\pm$ 1.7 $\pm$ 2.1	76	<sup>9</sup> ACCIARRI	99T L3	$e^+e^- \rightarrow e^+e^-\eta_C$	
27 $\pm$ 16 $\pm$ 10	5	<sup>5</sup> SHIRAI	98 AMY	58 $e^+e^-$	
6.7 $\pm$ 2.4		<sup>4</sup> ARMSTRONG	95F E760	$\bar{p}p \rightarrow \gamma\gamma$	
11.3 $\pm$ 4.2		<sup>10</sup> ALBRECHT	94H ARG	$e^+e^- \rightarrow e^+e^-\eta_C$	
8.0 $\pm$ 2.3 $\pm$ 2.4	17	<sup>11</sup> ADRIANI	93N L3	$e^+e^- \rightarrow e^+e^-\eta_C$	
5.9 $\pm$ 2.1		<sup>7</sup> CHEN	90B CLEO	$e^+e^- \rightarrow e^+e^-\eta_C$	
6.4 $\pm$ 5.0		<sup>12</sup> AIHARA	88D TPC	$e^+e^- \rightarrow e^+e^-\chi$	
4.3 $\pm$ 3.4		<sup>4</sup> BAGLIN	87B SPEC	$\bar{p}p \rightarrow \gamma\gamma$	
28 $\pm$ 15		<sup>5,13</sup> BERGER	86 PLUT	$\gamma\gamma \rightarrow K\bar{K}\pi$	
<sup>1</sup> Assuming there is no interference with the non-resonant background.					
<sup>2</sup> Calculated by us using $\Gamma(\eta_C \rightarrow K\bar{K}\pi) \times \Gamma(\eta_C \rightarrow \gamma\gamma) / \Gamma = 0.44 \pm 0.05$ keV from PDG 06 and $B(\eta_C \rightarrow K\bar{K}\pi) = (8.5 \pm 1.8)\%$ from AUBERT 06E.					
<sup>3</sup> Systematic errors not evaluated.					
<sup>4</sup> Normalized to $B(\eta_C \rightarrow p\bar{p}) = (1.3 \pm 0.4) \times 10^{-3}$ .					
<sup>5</sup> Normalized to $B(\eta_C \rightarrow K^\pm K_S^0 \pi^\mp)$ .					
<sup>6</sup> Average of $K_S^0 K^\pm \pi^\mp$ , $\pi^+\pi^-K^+K^-$ , and $2(K^+K^-)$ decay modes.					
<sup>7</sup> Normalized to the sum of $B(\eta_C \rightarrow K^\pm K_S^0 \pi^\mp)$ , $B(\eta_C \rightarrow K^+K^-\pi^+\pi^-)$ , and $B(\eta_C \rightarrow 2\pi^+2\pi^-)$ .					
<sup>8</sup> Superseded by ASNER 04.					
<sup>9</sup> Normalized to the sum of 9 branching ratios.					
<sup>10</sup> Normalized to the sum of $B(\eta_C \rightarrow K^\pm K_S^0 \pi^\mp)$ , $B(\eta_C \rightarrow \phi\phi)$ , $B(\eta_C \rightarrow K^+K^-\pi^+\pi^-)$ , and $B(\eta_C \rightarrow 2\pi^+2\pi^-)$ .					
<sup>11</sup> Superseded by ACCIARRI 99T.					
<sup>12</sup> Normalized to the sum of $B(\eta_C \rightarrow K^\pm K_S^0 \pi^\mp)$ , $B(\eta_C \rightarrow 2K^+2K^-)$ , $B(\eta_C \rightarrow K^+K^-\pi^+\pi^-)$ , and $B(\eta_C \rightarrow 2\pi^+2\pi^-)$ .					
<sup>13</sup> Re-evaluated by AIHARA 88D.					

$\frac{\Gamma(\eta'(958)\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}}{\text{VALUE (keV)}}$	$\frac{\text{EVTS}}{\text{DOCUMENT ID}}$	$\frac{\text{TECN}}{\text{COMMENT}}$	$\frac{\Gamma_{47}/\Gamma}{\text{COMMENT}}$
$75.8^{+6.3}_{-6.2} \pm 8.4$	486	<sup>1</sup> ZHANG	12A BELL $e^+e^- \rightarrow \eta' \pi^+ \pi^-$

<sup>1</sup> Assuming there is no interference with the non-resonant background.

$\Gamma(K^*(892)\bar{K}^*(892)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{4\gamma}/\Gamma$
VALUE (eV)		EVTS	DOCUMENT ID	TECN	COMMENT
<b>35 ± 6</b>	<b>OUR FIT</b>				
<b>32.4 ± 4.2 ± 5.8</b>		882 ± 115	UEHARA	08 BELL	$\gamma\gamma \rightarrow \pi^+\pi^-K^+K^-$

$7.75 \pm 0.66 \pm 0.62$        $386 \pm 31$        $^1 \text{LIU}$       12B BELL       $\gamma\gamma \rightarrow 2(K^+K^-)$   
 • • • We do not use the following data for averages, fits, limits, etc. • • •  
 $6.8 \pm 1.2 \pm 1.3$        $132 \pm 23$       UEHARA      08 BELL       $\gamma\gamma \rightarrow 2(K^+K^-)$   
 $^1$ Supersedes UEHARA 08. Using  $\text{B}(\phi \rightarrow K^+K^-) = (48.9 \pm 0.5)\%$ .

<sup>1</sup> Using  $\mathcal{B}(\omega \rightarrow \pi^+ \pi^- \pi^0) = (89.2 \pm 0.7)\%$ .

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.49	90	<sup>1</sup> LIU	12B	BELL	$\gamma\gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$
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<sup>1</sup> Using  $B(\phi \rightarrow K^+ K^-) = (48.9 \pm 0.5)\%$  and  $B(\omega \rightarrow \pi^+ \pi^- \pi^0) = (89.2 \pm 0.7)\%$ .

$50 \pm 13$ OUR FIT				
$69 \pm 17 \pm 12$	$3182 \pm 766$	UEHARA	08	BELL $\gamma\gamma \rightarrow 2(\pi^+ \pi^-)$

<b><math>49 \pm 9 \pm 13</math></b>	$1128 \pm 206$	UEHARA	08	BELL	$\gamma\gamma \rightarrow \pi^+ \pi^- K^+ K^-$
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0.398 ± 0.021 OUR FIT							
0.407 ± 0.027 OUR AVERAGE		Error includes scale factor of 1.2.					
0.374 ± 0.009 ± 0.031	14k	<sup>1</sup> LEES	10	BABR	$10.6 e^+ e^- \rightarrow K_S^0 K^\pm \pi^\mp$		
0.407 ± 0.022 ± 0.028		<sup>2,3</sup> ASNER	04	CLEO	$\gamma\gamma \rightarrow \eta_c \rightarrow K_S^0 K^\pm \pi^\mp$		
0.60 ± 0.12 ± 0.09	41	<sup>3,4</sup> ABDALLAH	03J	DLPH	$\gamma\gamma \rightarrow K_S^0 K^\pm \pi^\mp$		
1.47 ± 0.87 ± 0.27		<sup>3</sup> SHIRAI	98	AMY	$\gamma\gamma \rightarrow \eta_c \rightarrow K^\pm K_S^0 \pi^\mp$		
0.84 ± 0.21		<sup>3</sup> ALBRECHT	94H	ARG	$\gamma\gamma \rightarrow K^\pm K_S^0 \pi^\mp$		
0.60 $\begin{smallmatrix} +0.23 \\ -0.20 \end{smallmatrix}$		<sup>3</sup> CHEN	90B	CLEO	$\gamma\gamma \rightarrow \eta_c K^\pm K_S^0 \pi^\mp$		
1.06 ± 0.41 ± 0.27	11	<sup>3</sup> BRAUNSCH...	89	TASS	$\gamma\gamma \rightarrow K \bar{K} \pi$		
1.5 $\begin{smallmatrix} +0.60 \\ -0.45 \end{smallmatrix}$ ± 0.3	7	<sup>3</sup> BERGER	86	PLUT	$\gamma\gamma \rightarrow K \bar{K} \pi$		
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●							
0.386 ± 0.008 ± 0.021	12k	<sup>5</sup> DEL-AMO-SA...	11M	BABR	$\gamma\gamma \rightarrow K_S^0 K^\pm \pi^\mp$		
0.418 ± 0.044 ± 0.022		<sup>3,6</sup> BRANDENB...	00B	CLE2	$\gamma\gamma \rightarrow \eta_c \rightarrow K^\pm K_S^0 \pi^\mp$		
< 0.63	95	<sup>3</sup> BEHREND	89	CELL	$\gamma\gamma \rightarrow K_S^0 K^\pm \pi^\mp$		
< 4.4	95	ALTHOFF	85B	TASS	$\gamma\gamma \rightarrow K \bar{K} \pi$		

$\Gamma(K^+K^-\pi^+\pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$				$\Gamma_{3247}/\Gamma$		
VALUE (eV)	EVTS			DOCUMENT ID	TECN	COMMENT
<b>35 ± 5</b>	<b>OUR FIT</b>					
<b>27 ± 6</b>	<b>OUR AVERAGE</b>					
25.7 ± 3.2 ± 4.9	2019 ± 248			UEHARA	08 BELL	$\gamma\gamma \rightarrow \pi^+\pi^-K^+K^-$
280 ± 100 ± 60	42			<sup>1</sup> ABDALLAH	03J DLPH	$\gamma\gamma \rightarrow \pi^+\pi^-K^+K^-$
170 ± 80 ± 20	13.9 ± 6.6			ALBRECHT	94H ARG	$\gamma\gamma \rightarrow \pi^+\pi^-K^+K^-$

$\Gamma(K^+K^-\pi^+\pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$
VALUE (keV)	EVS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.190 ± 0.006 ± 0.028	11k	<sup>1</sup> DEL-AMO-SA...11M	BABR	$\gamma\gamma \rightarrow K^+K^-\pi^+\pi^-$	
<sup>1</sup> Not independent from other measurements reported in DEL-AMO-SANCHEZ 11M.					



# Meson Particle Listings

## $\eta_c(1S)$

$\Gamma(f_2(1270)f_2(1270))/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.98 \pm 0.25</math> OUR FIT</b>					

<b><math>0.77^{+0.25}_{-0.30} \pm 0.17</math></b>	$91.2 \pm 19.8$	<sup>1</sup> ABLIKIM	04M BES	$J/\psi \rightarrow \gamma 2\pi^+ 2\pi^-$	
<sup>1</sup> ABLIKIM 04M reports $[\Gamma(\eta_c(1S) \rightarrow f_2(1270)f_2(1270))/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] = (1.3 \pm 0.3^{+0.3}_{-0.4}) \times 10^{-4}$ which we divide by our best value $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(f_0(980)\eta)/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \eta$	

$\Gamma(f_0(1500)\eta)/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \eta$	

$\Gamma(f_0(2200)\eta)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \eta$	

$\Gamma(a_0(980)\pi)/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \pi^0$	

$\Gamma(a_0(1320)\pi)/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \pi^0$	

$\Gamma(a_0(1450)\pi)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \pi^0$	

$\Gamma(a_0(1950)\pi)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
seen	12k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$	

<sup>1</sup> From a model-independent partial wave analysis.

$\Gamma(a_2(1950)\pi)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
not seen	12k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$	

<sup>1</sup> From a model-independent partial wave analysis assuming the existence of a hypothetical tensor isovector  $a_2(1950)$ .

$\Gamma(K_S^*(1430)\overline{K})/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
seen	12k	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$	
seen		LEES	14E	BABR Dalitz anal. of $\eta_c \rightarrow K^+ K^- \eta/\pi^0$	

<sup>1</sup> From a model-independent partial wave analysis.

$\Gamma(K_S^*(1430)\overline{K})/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	LEES	14E	BABR	Dalitz anal. of $\eta_c \rightarrow K^+ K^- \pi^0$	

$\Gamma(K_S^*(1950)\overline{K})/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
seen	12K	<sup>1</sup> LEES	16A	BABR $\gamma\gamma \rightarrow \eta_c(1S) \rightarrow K\overline{K}\pi$	
seen		LEES	14E	BABR Dalitz anal. of $\eta_c \rightarrow K^+ K^- \eta/\pi^0$	

<sup>1</sup> From a Dalitz plot analysis using an isobar model.

$\Gamma(K\overline{K}\pi)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>7.3 \pm 0.5</math> OUR FIT</b>					

<b><math>6.5 \pm 0.6</math> OUR AVERAGE</b>					
$6.3 \pm 1.3 \pm 0.6$	55	<sup>1,2</sup> ABLIKIM	12N	BES3 $\psi(2S) \rightarrow \pi^0 \gamma K^+ K^- \pi^0$	
$7.9 \pm 1.4 \pm 0.7$	107	<sup>3,4</sup> ABLIKIM	12N	BES3 $\psi(2S) \rightarrow \pi^0 \gamma K_S^0 K^\mp \pi^\pm$	
$8.5 \pm 1.8$		<sup>5</sup> AUBERT	06E	BABR $B^\pm \rightarrow K^\pm X_{C\overline{C}}^\mp$	
$5.1 \pm 2.1$	0.6k	<sup>6</sup> BAI	04	BES $J/\psi \rightarrow \gamma K^\pm \pi^\mp K_S^0$	
$6.90 \pm 1.42 \pm 1.32$	33	<sup>6</sup> BISELLO	91	DM2 $J/\psi \rightarrow \gamma K^\pm K^- \pi^0$	
$5.43 \pm 0.94 \pm 0.94$	68	<sup>6</sup> BISELLO	91	DM2 $J/\psi \rightarrow \gamma K^\pm \pi^\mp K_S^0$	
$4.8 \pm 1.7$	95	<sup>6,7</sup> BALTRUSAIT..86	MRK3	$J/\psi \rightarrow \eta_c \gamma$	
$16.1^{+9.2}_{-7.3}$		<sup>8,9</sup> HIMEL	80B	MRK2 $\psi(2S) \rightarrow \eta_c \gamma$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 10.7    90% CL    <sup>6,10</sup> PARTRIDGE    80B    CBAL     $J/\psi \rightarrow \eta_c \gamma$

- <sup>1</sup> ABLIKIM 12N quotes  $B(\psi(2S) \rightarrow \pi^0 h_c) \cdot B(h_c \rightarrow \gamma \eta_c) \cdot B(\eta_c \rightarrow K^+ K^- \pi^0) = (4.54 \pm 0.76 \pm 0.48) \times 10^{-6}$  which we multiply by 6 to account for isospin symmetry.
- <sup>2</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow K\overline{K}\pi)/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (27.24 \pm 4.56 \pm 2.88) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>3</sup> ABLIKIM 12N quotes  $B(\psi(2S) \rightarrow \pi^0 h_c) \cdot B(h_c \rightarrow \gamma \eta_c) \cdot B(\eta_c \rightarrow K_S^0 K^\pm \pi^\mp) = (11.35 \pm 1.25 \pm 1.50) \times 10^{-6}$  which we multiply by 3 to account for isospin symmetry.
- <sup>4</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow K\overline{K}\pi)/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (34.05 \pm 3.75 \pm 4.50) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>5</sup> Determined from the ratio of  $B(B^\pm \rightarrow K^\pm \eta_c) B(\eta_c \rightarrow K\overline{K}\pi) = (7.4 \pm 0.5 \pm 0.7) \times 10^{-5}$  reported in AUBERT,B 04b and  $B(B^\pm \rightarrow K^\pm \eta_c) = (8.7 \pm 1.5) \times 10^{-3}$  reported in AUBERT 06e.
- <sup>6</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = 0.0127 \pm 0.0036$ . Where relevant, the error in this branching ratio is treated as a common systematic in computing averages.
- <sup>7</sup> Average from  $K^+ K^- \pi^0$  and  $K^\pm K_S^0 \pi^\mp$  decay channels.
- <sup>8</sup>  $K^\pm K_S^0 \pi^\mp$  corrected to  $K\overline{K}\pi$  by factor 3. KS, MR.
- <sup>9</sup> Estimated using  $B(\psi(2S) \rightarrow \gamma \eta_c(1S)) = 0.0028 \pm 0.0006$ .
- <sup>10</sup>  $K^+ K^- \pi^0$  corrected to  $K\overline{K}\pi$  by factor 6. KS, MR

$\Gamma(\phi K^+ K^-)/\Gamma(K\overline{K}\pi)$					$\Gamma_6/\Gamma_{28}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.052^{+0.016}_{-0.014} \pm 0.014</math></b>	7	<sup>1</sup> HUANG	03	BELL $B^\pm \rightarrow K^\pm \phi \phi$	

<sup>1</sup> Using  $B(B^+ \rightarrow \eta_c K^+) = (1.25 \pm 0.12^{+0.10}_{-0.12}) \times 10^{-3}$  from FANG 03 and  $B(\eta_c \rightarrow K\overline{K}\pi) = (5.5 \pm 1.7) \times 10^{-2}$ .

$\Gamma(K\overline{K}\eta)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-2}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.35 \pm 0.16</math> OUR FIT</b>					
<b><math>1.0 \pm 0.5 \pm 0.2</math></b>	7	<sup>1,2</sup> ABLIKIM	12N	BES3 $\psi(2S) \rightarrow \pi^0 \gamma K^+ K^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<3.1	90	<sup>3</sup> BALTRUSAIT..86	MRK3	$J/\psi \rightarrow \eta_c \gamma$	

- <sup>1</sup> ABLIKIM 12N quotes  $B(\psi(2S) \rightarrow \pi^0 h_c) \cdot B(h_c \rightarrow \gamma \eta_c) \cdot B(\eta_c \rightarrow K^+ K^- \eta) = (2.11 \pm 1.01 \pm 0.32) \times 10^{-6}$  which we multiply by 2 to account for isospin symmetry.
- <sup>2</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow K\overline{K}\eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))] \times [B(h_c(1P) \rightarrow \eta_c(1S)\gamma)] = (4.22 \pm 2.02 \pm 0.64) \times 10^{-6}$  which we divide by our best values  $B(\psi(2S) \rightarrow \pi^0 h_c(1P)) = (8.6 \pm 1.3) \times 10^{-4}$ ,  $B(h_c(1P) \rightarrow \eta_c(1S)\gamma) = (51 \pm 6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.
- <sup>3</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = 0.0127 \pm 0.0036$ .

$\Gamma(K\overline{K}\eta)/\Gamma(K\overline{K}\pi)$					$\Gamma_{29}/\Gamma_{28}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.186 \pm 0.018</math> OUR FIT</b>					
<b><math>0.190 \pm 0.008 \pm 0.017</math></b>	5.4k	<sup>1</sup> LEES	14E	BABR $\gamma\gamma \rightarrow K^+ K^- \eta/\pi^0$	

<sup>1</sup> LEES 14E reports  $B(\eta_c(1S) \rightarrow K^+ K^- \eta)/B(\eta_c(1S) \rightarrow K^+ K^- \pi^0) = 0.571 \pm 0.025 \pm 0.051$ , which we divide by 3 to account for isospin symmetry. It uses both  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decays.

$\Gamma(\eta\pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.7 \pm 0.4 \pm 0.1</math></b>	33	<sup>1</sup> ABLIKIM	12N	BES3 $\psi(2S) \rightarrow \pi^0 \gamma \eta \pi^+ \pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$5.4 \pm 2.0$	75	<sup>2</sup> BALTRUSAIT..86	MRK3	$J/\psi \rightarrow \eta_c \gamma$	
$3.7 \pm 1.3 \pm 2.0$	18	<sup>2</sup> PARTRIDGE	80B	CBAL $J/\psi \rightarrow \eta \pi^+ \pi^- \gamma$	

- <sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (7.22 \pm 1.47 \pm 1.11) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = 0.0127 \pm 0.0036$ . Where relevant, the error in this branching ratio is treated as a common systematic in computing averages.

$\Gamma(\eta(2\pi^+ \pi^-))/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>4.4 \pm 1.2 \pm 0.4</math></b>	39	<sup>1</sup> ABLIKIM	12N	BES3 $\psi(2S) \rightarrow \pi^0 \gamma \eta 2(\pi^+ \pi^-)$	

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow \eta 2(\pi^+ \pi^-))/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (19.17 \pm 3.77 \pm 3.72) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.



$\Gamma(K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{32}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>6.9 ± 1.1 OUR FIT</b>			
<b>11.2 ± 1.9 OUR AVERAGE</b>			
9.7 ± 2.2 ± 0.9	38	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma K^+ K^- \pi^+ \pi^-$
12 ± 4	0.4k	<sup>2</sup> BAI	04 BES $J/\psi \rightarrow \gamma K^+ K^- \pi^+ \pi^-$
21 ± 7	110	<sup>2</sup> BALTRUSAIT...86	MRK3 $J/\psi \rightarrow \eta_c \gamma$
14 +22 -9		<sup>3</sup> HIMEL	80B MRK2 $\psi(2S) \rightarrow \eta_c \gamma$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (4.16 \pm 0.76 \pm 0.59) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = 0.0127 \pm 0.0036$ . Where relevant, the error in this branching ratio is treated as a common systematic in computing averages.

<sup>3</sup> Estimated using  $B(\psi(2S) \rightarrow \gamma \eta_c(1S)) = 0.0028 \pm 0.0006$ .

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma(K\bar{K}\pi)$		$\Gamma_{33}/\Gamma_{28}$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.477 ± 0.017 ± 0.070</b>	11k	<sup>1</sup> DEL-AMO-SA...11M	BABR $\gamma\gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$

<sup>1</sup> We have multiplied the value of  $\Gamma(K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma(K_S^0 K^\pm \pi^\mp)$  reported in DEL-AMO-SANCHEZ 11M by a factor 1/3 to obtain  $\Gamma(K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma(K\bar{K}\pi)$ . Not independent from other measurements reported in DEL-AMO-SANCHEZ 11M.

$\Gamma(K^0 K^- \pi^+ \pi^- \pi^+ + \text{c.c.})/\Gamma_{\text{total}}$		$\Gamma_{34}/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>5.6 ± 1.4 ± 0.5</b>	43	<sup>1,2</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma K_S^0 K^\mp \pi^\mp 2\pi^\pm$

<sup>1</sup> ABLIKIM 12N quotes  $B(\psi(2S) \rightarrow \pi^0 h_c) \cdot B(h_c \rightarrow \gamma \eta_c) \cdot B(\eta_c \rightarrow K_S^0 K^- \pi^- 2\pi^+)$  =  $(12.01 \pm 2.22 \pm 2.04) \times 10^{-6}$  which we multiply by 2 to take c.c. into account.

<sup>2</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow K^0 K^- \pi^+ \pi^- \pi^+ + \text{c.c.})/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (24.02 \pm 4.44 \pm 4.08) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+ K^- 2(\pi^+ \pi^-))/\Gamma_{\text{total}}$		$\Gamma_{35}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>7.5 ± 2.4 OUR AVERAGE</b>			
8 ± 4 ± 1	10	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma K^+ K^- 2(\pi^+ \pi^-)$
7.2 ± 2.4 ± 1.6	100	<sup>2</sup> ABLIKIM	06A BES2 $J/\psi \rightarrow K^+ K^- 2(\pi^+ \pi^-) \gamma$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow K^+ K^- 2(\pi^+ \pi^-))/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (3.60 \pm 1.71 \pm 0.64) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ABLIKIM 06A reports  $[\Gamma(\eta_c(1S) \rightarrow K^+ K^- 2(\pi^+ \pi^-))/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma \eta_c(1S))] = (1.21 \pm 0.32 \pm 0.24) \times 10^{-4}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2(K^+ K^-))/\Gamma_{\text{total}}$		$\Gamma_{36}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>1.46 ± 0.30 OUR FIT</b>			
<b>2.2 ± 0.9 ± 0.2</b>	7	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma 2(K^+ K^-)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.4 +0.5 -0.4	± 0.6	14.5 +4.6 -3.0	<sup>2</sup> HUANG	03 BELL	$B^+ \rightarrow 2(K^+ K^-) K^+$
21 ± 10	± 6		<sup>3</sup> ALBRECHT	94H ARG	$\gamma\gamma \rightarrow K^+ K^- K^+ K^-$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow 2(K^+ K^-))/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (0.94 \pm 0.37 \pm 0.14) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Using  $B(B^+ \rightarrow \eta_c K^+) = (1.25 \pm 0.12 \pm 0.10) \times 10^{-3}$  from FANG 03 and  $B(\eta_c \rightarrow K\bar{K}\pi) = (5.5 \pm 1.7) \times 10^{-2}$ .

<sup>3</sup> Normalized to the sum of  $B(\eta_c \rightarrow K^\pm K_S^0 \pi^\mp)$ ,  $B(\eta_c \rightarrow \phi \phi)$ ,  $B(\eta_c \rightarrow K^+ K^- \pi^+ \pi^-)$ , and  $B(\eta_c \rightarrow 2\pi^+ 2\pi^-)$ .

$\Gamma(2(K^+ K^-))/\Gamma(K\bar{K}\pi)$		$\Gamma_{36}/\Gamma_{28}$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.020 ± 0.004 OUR FIT</b>			
<b>0.024 ± 0.007 OUR AVERAGE</b>			
0.023 ± 0.007 ± 0.006		AUBERT,B	04B BABR $B^\pm \rightarrow K^\pm \eta_c$
0.026 +0.009 -0.007 ± 0.007	15	<sup>1</sup> HUANG	03 BELL $B^\pm \rightarrow K^\pm (2K^+ 2K^-)$

<sup>1</sup> Using  $B(B^+ \rightarrow \eta_c K^+) = (1.25 \pm 0.12 \pm 0.10) \times 10^{-3}$  from FANG 03 and  $B(\eta_c \rightarrow K\bar{K}\pi) = (5.5 \pm 1.7) \times 10^{-2}$ .

$\Gamma(\pi^+ \pi^- \pi^0 \pi^0)/\Gamma_{\text{total}}$		$\Gamma_{37}/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>4.7 ± 0.9 ± 0.4</b>	118	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma \pi^+ \pi^- 2\pi^0$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow \pi^+ \pi^- \pi^0 \pi^0)/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (20.31 \pm 2.20 \pm 3.33) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2(\pi^+ \pi^-))/\Gamma_{\text{total}}$		$\Gamma_{38}/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.97 ± 0.12 OUR FIT</b>			
<b>1.35 ± 0.21 OUR AVERAGE</b>			
1.74 ± 0.32 ± 0.15	100	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma 2(\pi^+ \pi^-)$
1.0 ± 0.5	542 ± 75	<sup>2</sup> BAI	04 BES $J/\psi \rightarrow \gamma 2(\pi^+ \pi^-)$
1.05 ± 0.17 ± 0.34	137	<sup>2</sup> BISELLO	91 DM2 $J/\psi \rightarrow \gamma 2\pi^+ 2\pi^-$
1.3 ± 0.6	25	<sup>2</sup> BALTRUSAIT...86	MRK3 $J/\psi \rightarrow \eta_c \gamma$
2.0 +1.5 -1.0		<sup>3</sup> HIMEL	80B MRK2 $\psi(2S) \rightarrow \eta_c \gamma$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow 2(\pi^+ \pi^-))/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (7.51 \pm 0.85 \pm 1.11) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = 0.0127 \pm 0.0036$ . Where relevant, the error in this branching ratio is treated as a common systematic in computing averages.

<sup>3</sup> Estimated using  $B(\psi(2S) \rightarrow \gamma \eta_c(1S)) = 0.0028 \pm 0.0006$ .

$\Gamma(2(\pi^+ \pi^- \pi^0))/\Gamma_{\text{total}}$		$\Gamma_{39}/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>17.4 ± 2.9 ± 1.5</b>	175	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma 2(\pi^+ \pi^- \pi^0)$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow 2(\pi^+ \pi^- \pi^0))/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (75.13 \pm 7.42 \pm 9.99) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3(\pi^+ \pi^-))/\Gamma_{\text{total}}$		$\Gamma_{40}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>18 ± 4 OUR AVERAGE</b>			
20 ± 5 ± 2	51	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma 3(\pi^+ \pi^-)$
15.3 ± 3.4 ± 3.3	479	<sup>2</sup> ABLIKIM	06A BES2 $J/\psi \rightarrow 3(\pi^+ \pi^-) \gamma$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow 3(\pi^+ \pi^-))/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (8.82 \pm 1.57 \pm 1.59) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ABLIKIM 06A reports  $[\Gamma(\eta_c(1S) \rightarrow 3(\pi^+ \pi^-))/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma \eta_c(1S))] = (2.59 \pm 0.32 \pm 0.47) \times 10^{-4}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p\bar{p})/\Gamma_{\text{total}}$		$\Gamma_{41}/\Gamma$	
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>15.0 ± 1.6 OUR FIT</b>			
<b>13.2 ± 2.7 OUR AVERAGE</b>			
15 ± 5 ± 1	15	<sup>1</sup> ABLIKIM	12N BES3 $\psi(2S) \rightarrow \pi^0 \gamma p\bar{p}$
15 ± 6	213 ± 33	<sup>2</sup> BAI	04 BES $J/\psi \rightarrow \gamma p\bar{p}$
10 ± 3 ± 4	18	<sup>2</sup> BISELLO	91 DM2 $J/\psi \rightarrow \gamma p\bar{p}$
11 ± 6	23	<sup>2</sup> BALTRUSAIT...86	MRK3 $J/\psi \rightarrow \eta_c \gamma$
29 +29 -15		<sup>3</sup> HIMEL	80B MRK2 $\psi(2S) \rightarrow \eta_c \gamma$
14.8 +2.0 +1.7 -2.4 -1.8	195	<sup>4</sup> WU	06 BELL $B^+ \rightarrow p\bar{p} K^+$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow p\bar{p})/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (0.65 \pm 0.19 \pm 0.10) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma \eta_c(1S)) = 0.0127 \pm 0.0036$ . Where relevant, the error in this branching ratio is treated as a common systematic in computing averages.

<sup>3</sup> Estimated using  $B(\psi(2S) \rightarrow \gamma \eta_c(1S)) = 0.0028 \pm 0.0006$ .

<sup>4</sup> WU 06 reports  $[\Gamma(\eta_c(1S) \rightarrow p\bar{p})/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow \eta_c K^+)] = (1.42 \pm 0.11 \pm 0.16) \times 10^{-6}$  which we divide by our best value  $B(B^+ \rightarrow \eta_c K^+) = (9.6 \pm 1.1) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

## $\eta_c(1S)$

$\Gamma(\rho\overline{\rho})/\Gamma(K\overline{K}\pi)$					$\Gamma_{41}/\Gamma_{28}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.0207±0.0021 OUR FIT</b>					
<b>0.021 ±0.002 +0.004 -0.006</b>	195	1 WU	06	BELL	$B^\pm \rightarrow K^\pm \rho\overline{p}$

<sup>1</sup> Using  $B(B^+ \rightarrow \eta_c K^+) = (1.25 \pm 0.12^{+0.10}_{-0.12}) \times 10^{-3}$  from FANG 03 and  $B(\eta_c \rightarrow K\overline{K}\pi) = (5.5 \pm 1.7) \times 10^{-2}$ .

$\Gamma(\rho\overline{\rho})/\Gamma_{\text{total}} \times \Gamma(\phi\phi)/\Gamma_{\text{total}}$					$\Gamma_{41}/\Gamma \times \Gamma_7/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT		
<b>0.26±0.05 OUR FIT</b>					
<b>4.0 +3.5 -3.2</b>	BAGLIN	89	SPEC	$\overline{p}p \rightarrow K^+ K^- K^+ K^-$	

$\Gamma(\rho\overline{p}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{42}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.36±0.13±0.03</b>					
	14	1 ABLIKIM	12N	BES3	$\psi(2S) \rightarrow \pi^0 \gamma \rho\overline{p} \pi^0$
<sup>1</sup> ABLIKIM 12N reports $[\Gamma(\eta_c(1S) \rightarrow \rho\overline{p}\pi^0)/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (1.53 \pm 0.49 \pm 0.23) \times 10^{-6}$ which we divide by our best value $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\Lambda\overline{\Lambda})/\Gamma_{\text{total}}$					$\Gamma_{43}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10.9±2.4 OUR FIT</b>					
<b>11.7±2.3±2.6</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
9.9 +2.7 -2.6 ±1.2	20	2 WU	06	BELL	$B^+ \rightarrow \Lambda\overline{\Lambda} K^+$
<20	90	3 BISELLO	91	DM2	$e^+ e^- \rightarrow \gamma \Lambda\overline{\Lambda}$

<sup>1</sup> ABLIKIM 12B reports  $[\Gamma(\eta_c(1S) \rightarrow \Lambda\overline{\Lambda})/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] = (0.198 \pm 0.021 \pm 0.032) \times 10^{-4}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> WU 06 reports  $[\Gamma(\eta_c(1S) \rightarrow \Lambda\overline{\Lambda})/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow \eta_c K^+)] = (0.95^{+0.25+0.08}_{-0.22-0.11}) \times 10^{-6}$  which we divide by our best value  $B(B^+ \rightarrow \eta_c K^+) = (9.6 \pm 1.1) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 0.0127 \pm 0.0036$ .

$\Gamma(\Lambda\overline{\Lambda})/\Gamma(\rho\overline{p})$					$\Gamma_{43}/\Gamma_{41}$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.72±0.16 OUR FIT</b>					
<b>0.67 +0.19 -0.16 ±0.12</b>	1 WU	06	BELL	$B^+ \rightarrow \rho\overline{p} K^+, \Lambda\overline{\Lambda} K^+$	

<sup>1</sup> Not independent from other  $\eta_c \rightarrow \Lambda\overline{\Lambda}, \rho\overline{p}$  branching ratios reported by WU 06.

$\Gamma(\Sigma^+ \Sigma^-)/\Gamma_{\text{total}}$					$\Gamma_{44}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.1±0.3±0.5</b>					
	112	1 ABLIKIM	13C	BES3	$J/\psi \rightarrow \gamma \rho\overline{p} \pi^0 \pi^0$
<sup>1</sup> ABLIKIM 13C reports $[\Gamma(\eta_c(1S) \rightarrow \Sigma^+ \Sigma^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] = (3.60 \pm 0.48 \pm 0.31) \times 10^{-5}$ which we divide by our best value $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\Xi^- \Xi^+)/\Gamma_{\text{total}}$					$\Gamma_{45}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.89±0.18±0.19</b>					
	78	1 ABLIKIM	13C	BES3	$J/\psi \rightarrow \gamma \Lambda\overline{\Lambda} \pi^+ \pi^-$
<sup>1</sup> ABLIKIM 13C reports $[\Gamma(\eta_c(1S) \rightarrow \Xi^- \Xi^+)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] = (1.51 \pm 0.27 \pm 0.14) \times 10^{-5}$ which we divide by our best value $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\pi^+ \pi^- \rho\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{46}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.3±1.7±0.5</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<12	90	HIMEL	80B	MRK2	$\psi(2S) \rightarrow \eta_c \gamma$

<sup>1</sup> ABLIKIM 12N reports  $[\Gamma(\eta_c(1S) \rightarrow \pi^+ \pi^- \rho\overline{p})/\Gamma_{\text{total}}] \times [\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}] = (2.30 \pm 0.65 \pm 0.36) \times 10^{-6}$  which we divide by our best value  $\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}} = (4.3 \pm 0.4) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{47}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.59±0.13 OUR FIT</b>					
<b>1.9 +0.7 -0.6 OUR AVERAGE</b>					

2.7 ±0.8 ±0.6	1 ABLIKIM	13i	BES3		
1.4 +0.7 -0.5 ±0.3	2 ADAMS	08	CLEO	$\psi(2S) \rightarrow \pi^+ \pi^- J/\psi$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.3 +1.0 -0.8 ±0.3	13	3 WICHT	08	BELL	$B^\pm \rightarrow K^\pm \gamma\gamma$
2.80 +0.67 -0.58 ±1.0		4 ARMSTRONG	95F	E760	$\overline{p}p \rightarrow \gamma\gamma$
<9	90	5 BISELLO	91	DM2	$J/\psi \rightarrow \gamma\gamma\gamma$
6 +4 -3 ±4		4 BAGLIN	87B	SPEC	$\overline{p}p \rightarrow \gamma\gamma$
<18	90	6 BLOOM	83	CBAL	$J/\psi \rightarrow \eta_c \gamma$

<sup>1</sup> ABLIKIM 13i reports  $[\Gamma(\eta_c(1S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] = (4.5 \pm 1.2 \pm 0.6) \times 10^{-6}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ADAMS 08 reports  $[\Gamma(\eta_c(1S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] = (2.4^{+1.1}_{-0.8} \pm 0.3) \times 10^{-6}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = (1.7 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> WICHT 08 reports  $[\Gamma(\eta_c(1S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow \eta_c K^+)] = (2.2^{+0.9+0.4}_{-0.7-0.2}) \times 10^{-7}$  which we divide by our best value  $B(B^+ \rightarrow \eta_c K^+) = (9.6 \pm 1.1) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> Not independent from the values of the total and two-photon width quoted by the same experiment.

<sup>5</sup> The quoted branching ratios use  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 0.0127 \pm 0.0036$ .

<sup>6</sup> Using  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 0.0127 \pm 0.0036$ .

$\Gamma(\gamma\gamma)/\Gamma(K\overline{K}\pi)$					$\Gamma_{47}/\Gamma_{28}$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.19±0.29 OUR FIT</b>					
<b>3.2 +1.3 +0.8 -1.0 ±0.6</b>	13	1 WICHT	08	BELL	$B^\pm \rightarrow K^\pm \gamma\gamma$

<sup>1</sup> Using  $B(B^+ \rightarrow \eta_c K^+) = (1.25 \pm 0.12^{+0.10}_{-0.12}) \times 10^{-3}$  from FANG 03 and  $B(\eta_c \rightarrow K\overline{K}\pi) = (5.5 \pm 1.7) \times 10^{-2}$ .

$\Gamma(\rho\overline{p})/\Gamma_{\text{total}} \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{41}/\Gamma \times \Gamma_{47}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.240±0.024 OUR FIT</b>					
<b>0.26 ±0.05 OUR AVERAGE</b>					
Error includes scale factor of 1.4.					
0.224 +0.038 -0.037 ±0.020	190	AMBROGIANI	03	E835	$\overline{p}p \rightarrow \eta_c \rightarrow \gamma\gamma$
0.336 +0.080 -0.070		ARMSTRONG	95F	E760	$\overline{p}p \rightarrow \gamma\gamma$
0.68 +0.42 -0.31	12	BAGLIN	87B	SPEC	$\overline{p}p \rightarrow \gamma\gamma$

## Charge conjugation (C), Parity (P), Lepton family number (LF) violating modes

$\Gamma(\pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{48}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;11</b>					
	90	1 ABLIKIM	11G	BES3	$J/\psi \rightarrow \gamma \pi^+ \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<70	90	2 ABLIKIM	06B	BES2	$J/\psi \rightarrow \pi^+ \pi^- \gamma$

<sup>1</sup> ABLIKIM 11G reports  $[\Gamma(\eta_c(1S) \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] < 1.82 \times 10^{-6}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 1.7 \times 10^{-2}$ .

<sup>2</sup> ABLIKIM 06B reports  $[\Gamma(\eta_c(1S) \rightarrow \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] < 1.1 \times 10^{-5}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 1.7 \times 10^{-2}$ .

$\Gamma(\pi^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{49}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;4</b>					
	90	1 ABLIKIM	11G	BES3	$J/\psi \rightarrow \gamma \pi^0 \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<40	90	2 ABLIKIM	06B	BES2	$J/\psi \rightarrow \pi^0 \pi^0 \gamma$

<sup>1</sup> ABLIKIM 11G reports  $[\Gamma(\eta_c(1S) \rightarrow \pi^0 \pi^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] < 6.0 \times 10^{-7}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 1.7 \times 10^{-2}$ .

<sup>2</sup> ABLIKIM 06B reports  $[\Gamma(\eta_c(1S) \rightarrow \pi^0 \pi^0)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] < 0.71 \times 10^{-5}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 1.7 \times 10^{-2}$ .

$\Gamma(K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;60</b>					
	90	1 ABLIKIM	06B	BES2	$J/\psi \rightarrow K^+ K^- \gamma$

<sup>1</sup> ABLIKIM 06B reports  $[\Gamma(\eta_c(1S) \rightarrow K^+ K^-)/\Gamma_{\text{total}}] \times [B(J/\psi(1S) \rightarrow \gamma\eta_c(1S))] < 0.96 \times 10^{-5}$  which we divide by our best value  $B(J/\psi(1S) \rightarrow \gamma\eta_c(1S)) = 1.7 \times 10^{-2}$ .









See key on page 601

## Meson Particle Listings

 $J/\psi(1S)$  $\Gamma(2(\pi^+\pi^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{103}\Gamma_5/\Gamma$ 

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>20.4±0.9±0.4</b>		LEES	12E	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 2π <sup>+</sup> 2π <sup>-</sup> γ
• • • We do not use the following data for averages, fits, limits, etc. • • •				
19.5±1.4±1.3	270	<sup>1</sup> AUBERT	05D	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 2(π <sup>+</sup> π <sup>-</sup> )γ
<sup>1</sup> Superseded by LEES 12E.				

 $\Gamma(3(\pi^+\pi^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{104}\Gamma_5/\Gamma$ 

VALUE (10 <sup>-2</sup> keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.37±0.16±0.14</b>	496	AUBERT	06D	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 3(π <sup>+</sup> π <sup>-</sup> )γ

 $\Gamma(2(\pi^+\pi^-\pi^0)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{105}\Gamma_5/\Gamma$ 

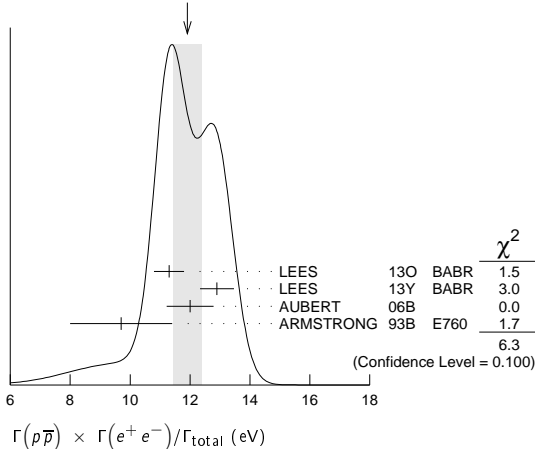
VALUE (10 <sup>-2</sup> keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.9±0.5±1.0</b>	761	AUBERT	06D	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 2(π <sup>+</sup> π <sup>-</sup> π <sup>0</sup> )γ

 $\Gamma(2(\pi^+\pi^-\eta)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{106}\Gamma_5/\Gamma$ 

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>13.1±2.4±0.1</b>	85	<sup>1</sup> AUBERT	07AU	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 2(π <sup>+</sup> π <sup>-</sup> )ηγ
<sup>1</sup> AUBERT 07AU reports [Γ(J/ψ(1S) → 2(π <sup>+</sup> π <sup>-</sup> )η) × Γ(J/ψ(1S) → e <sup>+</sup> e <sup>-</sup> )/Γ <sub>total</sub> ] × [B(η → 2γ)] = 5.16 ± 0.85 ± 0.39 eV which we divide by our best value B(η → 2γ) = (39.41 ± 0.20) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

 $\Gamma(p\bar{p}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{108}\Gamma_5/\Gamma$ 

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>11.9±0.5 OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.		
11.3±0.4±0.3	821	LEES	13O	BABR e <sup>+</sup> e <sup>-</sup> → p $\bar{p}$ γ
12.9±0.4±0.4	918	LEES	13Y	BABR e <sup>+</sup> e <sup>-</sup> → p $\bar{p}$ γ
12.0±0.6±0.5	438	AUBERT	06B	e <sup>+</sup> e <sup>-</sup> → p $\bar{p}$ γ
9.7±1.7		<sup>1</sup> ARMSTRONG	93B	E760 p $\bar{p}$ → e <sup>+</sup> e <sup>-</sup>
<sup>1</sup> Using Γ <sub>total</sub> = 85.5 + <sup>6.1</sup> <sub>-5.8</sub> MeV.				

WEIGHTED AVERAGE  
11.9±0.5 (Error scaled by 1.4) $\Gamma(\Sigma^0\bar{\Sigma}^0) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{121}\Gamma_5/\Gamma$ 

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b>6.4±1.2±0.6</b>	AUBERT	07BD	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → Σ <sup>0</sup> $\bar{\Sigma}^0$ γ

 $\Gamma(2(\pi^+\pi^-)K^+K^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{122}\Gamma_5/\Gamma$ 

VALUE (10 <sup>-2</sup> keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.75±0.23±0.17</b>	205	AUBERT	06D	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup> 2(π <sup>+</sup> π <sup>-</sup> )γ

 $\Gamma(\Lambda\bar{\Lambda}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{128}\Gamma_5/\Gamma$ 

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b>10.7±0.9±0.7</b>	AUBERT	07BD	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → Λ $\bar{\Lambda}$ γ

 $\Gamma(2(K^+K^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{131}\Gamma_5/\Gamma$ 

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.11±0.39±0.30</b>	156 ± 15	AUBERT	07AK	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 2(K <sup>+</sup> K <sup>-</sup> )γ
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.0 ± 0.7 ± 0.6	38	<sup>1</sup> AUBERT	05D	BABR 10.6 e <sup>+</sup> e <sup>-</sup> → 2(K <sup>+</sup> K <sup>-</sup> )γ
<sup>1</sup> Superseded by AUBERT 07AK.				

 $\Gamma(K^+K^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{133}\Gamma_5/\Gamma$ 

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.78±0.11±0.05	462	<sup>1</sup> LEES	15J	BABR e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup> γ
1.94±0.11±0.05	462	<sup>2</sup> LEES	15J	BABR e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup> γ
1.42±0.23±0.08	51	<sup>3</sup> LEES	13Q	BABR e <sup>+</sup> e <sup>-</sup> → K <sup>+</sup> K <sup>-</sup> γ
<sup>1</sup> sin φ > 0.				
<sup>2</sup> sin φ < 0.				
<sup>3</sup> Interference with non-resonant K <sup>+</sup> K <sup>-</sup> production not taken into account.				

 $J/\psi(1S)$  BRANCHING RATIOSFor the first four branching ratios, see also the partial widths, and (partial widths) × Γ(e<sup>+</sup>e<sup>-</sup>)/Γ<sub>total</sub> above. $\Gamma(\text{hadrons})/\Gamma_{\text{total}}$   $\Gamma_1/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.877±0.005 OUR AVERAGE</b>			
0.878±0.005	BAI	95B	BES e <sup>+</sup> e <sup>-</sup>
0.86 ± 0.02	BOYARSKI	75	MRK1 e <sup>+</sup> e <sup>-</sup>

 $\Gamma(\text{virtual}\gamma \rightarrow \text{hadrons})/\Gamma_{\text{total}}$   $\Gamma_2/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.135±0.003</b>	<sup>1,2</sup> SETH	04	RVUE e <sup>+</sup> e <sup>-</sup>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.17 ± 0.02	<sup>1</sup> BOYARSKI	75	MRK1 e <sup>+</sup> e <sup>-</sup>
<sup>1</sup> Included in Γ(hadrons)/Γ <sub>total</sub> .			
<sup>2</sup> Using B(J/ψ → ℓ <sup>+</sup> ℓ <sup>-</sup> ) = (5.90 ± 0.09)% from RPP-2002 and R = 2.28 ± 0.04 determined by a fit to data from BAI 00 and BAI 02C.			

 $\Gamma(g\bar{g}g)/\Gamma_{\text{total}}$   $\Gamma_3/\Gamma$ 

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>64.1±1.0</b>	6 M	<sup>1</sup> BESSON	08	CLEO ψ(2S) → π <sup>+</sup> π <sup>-</sup> + hadrons
<sup>1</sup> Calculated using the value Γ(γγg)/Γ(gḡg) = 0.137 ± 0.001 ± 0.016 ± 0.004 from BESSON 08 and the PDG 08 values of B(ℓ <sup>+</sup> ℓ <sup>-</sup> ), B(virtual γ → hadrons), and B(γη <sub>c</sub> ). The statistical error is negligible and the systematic error is partially correlated with that of Γ(γγg)/Γ <sub>total</sub> measurement of BESSON 08.				

 $\Gamma(\gamma\bar{g}g)/\Gamma_{\text{total}}$   $\Gamma_4/\Gamma$ 

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.79±1.05</b>	200 k	<sup>1</sup> BESSON	08	CLEO ψ(2S) → π <sup>+</sup> π <sup>-</sup> γ + hadrons
<sup>1</sup> Calculated using the value Γ(γγg)/Γ(gḡg) = 0.137 ± 0.001 ± 0.016 ± 0.004 from BESSON 08 and the value of Γ(gḡg)/Γ <sub>total</sub> . The statistical error is negligible and the systematic error is partially correlated with that of Γ(gḡg)/Γ <sub>total</sub> measurement of BESSON 08.				

 $\Gamma(\gamma\bar{g}g)/\Gamma(g\bar{g}g)$   $\Gamma_4/\Gamma_3$ 

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>13.7±0.1±0.7</b>	6 M	BESSON	08	CLEO ψ(2S) → π <sup>+</sup> π <sup>-</sup> J/ψ

 $\Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$ 

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.971±0.032 OUR AVERAGE</b>				
5.983±0.007±0.037	720k	ABLIKIM	13R	BES3 ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
5.945±0.067±0.042	15k	LI	05c	CLEO ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
5.90 ± 0.05 ± 0.10		BAI	98D	BES ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
6.09 ± 0.33		BAI	95B	BES e <sup>+</sup> e <sup>-</sup>
5.92 ± 0.15 ± 0.20		COFFMAN	92	MRK3 ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
6.9 ± 0.9		BOYARSKI	75	MRK1 e <sup>+</sup> e <sup>-</sup>

 $\Gamma(e^+e^-\gamma)/\Gamma_{\text{total}}$   $\Gamma_6/\Gamma$ 

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>8.8±1.3±0.4</b>	<sup>1</sup> ARMSTRONG	96	E760 p $\bar{p}$ → e <sup>+</sup> e <sup>-</sup> γ
<sup>1</sup> For E <sub>γ</sub> > 100 MeV.			

 $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$ 

VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.961±0.033 OUR AVERAGE</b>				
5.973±0.007±0.038	770k	ABLIKIM	13R	BES3 ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
5.960±0.065±0.050	17k	LI	05c	CLEO ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
5.84 ± 0.06 ± 0.10		BAI	98D	BES ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
6.08 ± 0.33		BAI	95B	BES e <sup>+</sup> e <sup>-</sup>
5.90 ± 0.15 ± 0.19		COFFMAN	92	MRK3 ψ(2S) → J/ψπ <sup>+</sup> π <sup>-</sup>
6.9 ± 0.9		BOYARSKI	75	MRK1 e <sup>+</sup> e <sup>-</sup>







## Meson Particle Listings

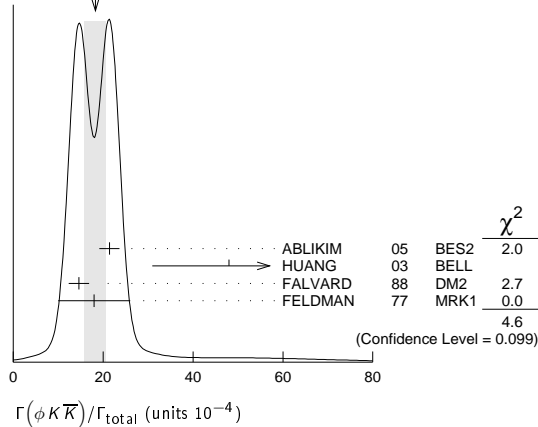
 $J/\psi(1S)$ 

<sup>1</sup> We have multiplied  $K^+ K^-$  measurement by 2 to obtain  $K\bar{K}$ .

<sup>2</sup> Using  $B(B^+ \rightarrow J/\psi K^+) = (1.01 \pm 0.05) \times 10^{-3}$ .

<sup>3</sup> Addition of  $\phi K^+ K^-$  and  $\phi K^0 \bar{K}^0$  branching ratios.

WEIGHTED AVERAGE  
18.3±2.4 (Error scaled by 1.5)

 $\Gamma(\phi f_0(1710) \rightarrow \phi K\bar{K})/\Gamma_{\text{total}}$  $\Gamma_{41}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.6±0.2±0.6</b>	1,2 FALVARD 88	DM2	$J/\psi \rightarrow \text{hadrons}$

<sup>1</sup> Including interference with  $f_2'(1525)$ .

<sup>2</sup> Includes unknown branching fraction  $f_0(1710) \rightarrow K\bar{K}$ .

 $\Gamma(\phi f_2(1270))/\Gamma_{\text{total}}$  $\Gamma_{42}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.72±0.13±0.02</b>	44 ± 7	1,2	AUBERT 07AK	BABR	$10.6 e^+ e^- \rightarrow \pi^+ \pi^- K^+ K^- \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.45	90		FALVARD 88	DM2	$J/\psi \rightarrow \text{hadrons}$
< 0.37	90		VANNUCCI 77	MRK1	$e^+ e^- \rightarrow \pi^+ \pi^- K^+ K^-$

<sup>1</sup> Using  $B(f_2(1270) \rightarrow \pi\pi) = (84.8^{+2.4}_{-1.2})\%$

<sup>2</sup> AUBERT 07AK reports  $[\Gamma(J/\psi(1S) \rightarrow \phi f_2(1270))/\Gamma_{\text{total}}] \times [\Gamma(J/\psi(1S) \rightarrow e^+ e^-)] = (4.02 \pm 0.65 \pm 0.33) \times 10^{-3}$  keV which we divide by our best value  $\Gamma(J/\psi(1S) \rightarrow e^+ e^-) = 5.55 \pm 0.14 \pm 0.02$  keV. Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\Delta(1232)^{++} \bar{\Delta}(1232)^{--})/\Gamma_{\text{total}}$  $\Gamma_{43}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.10±0.09±0.28</b>	233	EATON 84	MRK2	$e^+ e^-$

 $\Gamma(\Sigma(1385)^- \bar{\Sigma}(1385)^+ \text{ (or c.c.)})/\Gamma_{\text{total}}$  $\Gamma_{44}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.10±0.12 OUR AVERAGE</b>				
1.23±0.07±0.30	0.8k	ABLIKIM 12P	BES2	$J/\psi \rightarrow \Sigma(1385)^- \bar{\Sigma}(1385)^+$
1.50±0.08±0.38	1k	ABLIKIM 12P	BES2	$J/\psi \rightarrow \Sigma(1385)^+ \bar{\Sigma}(1385)^-$
1.00±0.04±0.21	0.6k	HENRARD 87	DM2	$e^+ e^- \rightarrow \Sigma^{*-}$
1.19±0.04±0.25	0.7k	HENRARD 87	DM2	$e^+ e^- \rightarrow \Sigma^{*+}$
0.86±0.18±0.22	56	EATON 84	MRK2	$e^+ e^- \rightarrow \Sigma^{*-}$
1.03±0.24±0.25	68	EATON 84	MRK2	$e^+ e^- \rightarrow \Sigma^{*+}$

 $\Gamma(\phi f_2'(1525))/\Gamma_{\text{total}}$  $\Gamma_{45}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8 ± 4 OUR AVERAGE</b>				Error includes scale factor of 2.7.
12.3±0.6±2.0		1,2 FALVARD 88	DM2	$J/\psi \rightarrow \text{hadrons}$
4.8±1.8	46	1 GIDAL 81	MRK2	$J/\psi \rightarrow K^+ K^- K^+ K^-$

<sup>1</sup> Re-evaluated using  $B(f_2'(1525) \rightarrow K\bar{K}) = 0.713$ .

<sup>2</sup> Including interference with  $f_0(1710)$ .

 $\Gamma(\phi \pi^+ \pi^-)/\Gamma_{\text{total}}$  $\Gamma_{46}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.94±0.09 OUR AVERAGE</b>				Error includes scale factor of 1.2.
0.96±0.13	103	1 AUBERT, BE 06D	BABR	$10.6 e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$
1.09±0.02±0.13		ABLIKIM 05	BES2	$J/\psi \rightarrow \phi \pi^+ \pi^-$
0.78±0.03±0.12		FALVARD 88	DM2	$J/\psi \rightarrow \text{hadrons}$
2.1 ± 0.9	23	FELDMAN 77	MRK1	$e^+ e^-$

<sup>1</sup> Derived by us. AUBERT, BE 06D measures  $\Gamma(J/\psi \rightarrow e^+ e^-) \times B(J/\psi \rightarrow \phi \pi^+ \pi^-) \times B(\phi \rightarrow K^+ K^-) = (2.61 \pm 0.30 \pm 0.18)$  eV

 $\Gamma(\phi \pi^0 \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{47}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.56±0.16</b>	23	1 AUBERT, BE 06D	BABR	$10.6 e^+ e^- \rightarrow K^+ K^- \pi^0 \pi^0 \gamma$

<sup>1</sup> Derived by us. AUBERT, BE 06D measures  $\Gamma(J/\psi \rightarrow e^+ e^-) \times B(J/\psi \rightarrow \phi \pi^0 \pi^0) \times B(\phi \rightarrow K^+ K^-) = (1.54 \pm 0.40 \pm 0.16)$  eV

 $\Gamma(\phi K^+ K_S^0 \pi^\mp)/\Gamma_{\text{total}}$  $\Gamma_{48}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.2±0.8 OUR AVERAGE</b>				
7.4±0.6±1.4	227 ± 19	ABLIKIM 08E	BES2	$e^+ e^- \rightarrow J/\psi$
7.4±0.9±1.1		FALVARD 88	DM2	$J/\psi \rightarrow \text{hadrons}$
7 ± 0.6±1.0	163 ± 15	BECKER 87	MRK3	$e^+ e^- \rightarrow \text{hadrons}$

 $\Gamma(\omega f_1(1420))/\Gamma_{\text{total}}$  $\Gamma_{49}/\Gamma$ 

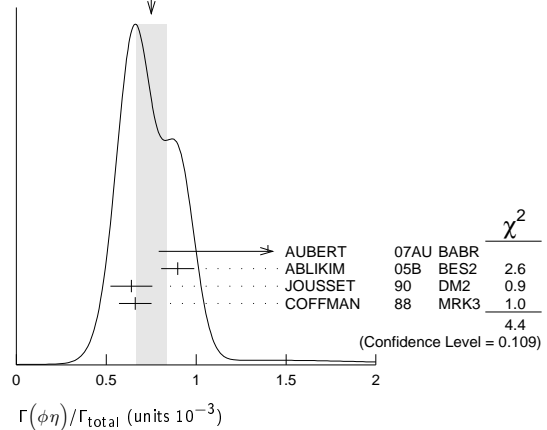
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.8±1.9±1.7</b>	111 ± <sup>31</sup> <sub>26</sub>	BECKER 87	MRK3	$e^+ e^- \rightarrow \text{hadrons}$

 $\Gamma(\phi \eta)/\Gamma_{\text{total}}$  $\Gamma_{50}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.75 ± 0.08 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
1.4 ± 0.6 ± 0.1	6	1 AUBERT 07AU	BABR	$10.6 e^+ e^- \rightarrow \phi \eta \gamma$
0.898±0.024±0.089		ABLIKIM 05B	BES2	$e^+ e^- \rightarrow J/\psi \rightarrow \text{hadr}$
0.64 ± 0.04 ± 0.11	346	JOUSSET 90	DM2	$J/\psi \rightarrow \text{hadrons}$
0.661±0.045±0.078		COFFMAN 88	MRK3	$e^+ e^- \rightarrow K^+ K^- \eta$

<sup>1</sup> AUBERT 07AU quotes  $\Gamma_{ee}^{J/\psi} \cdot B(J/\psi \rightarrow \phi \eta) \cdot B(\phi \rightarrow K^+ K^-) \cdot B(\eta \rightarrow \gamma \gamma) = 0.84 \pm 0.37 \pm 0.05$  eV.

WEIGHTED AVERAGE  
0.75±0.08 (Error scaled by 1.5)

 $\Gamma(\Xi^0 \Xi^0)/\Gamma_{\text{total}}$  $\Gamma_{51}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.20±0.12±0.21</b>	206	ABLIKIM 08o	BES2	$e^+ e^- \rightarrow J/\psi$

 $\Gamma(\Xi(1530)^- \Xi^+)/\Gamma_{\text{total}}$  $\Gamma_{52}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.59±0.09±0.12</b>	75 ± 11	HENRARD 87	DM2	$e^+ e^-$

 $\Gamma(\rho K^- \Sigma(1385)^0)/\Gamma_{\text{total}}$  $\Gamma_{53}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.51±0.26±0.18</b>	89	EATON 84	MRK2	$e^+ e^-$

 $\Gamma(\omega \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{54}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.45 ± 0.05 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
0.538±0.012±0.065	2090	1 ABLIKIM 06F	BES2	$J/\psi \rightarrow \omega \pi^0$
0.360±0.028±0.054	222	JOUSSET 90	DM2	$J/\psi \rightarrow \text{hadrons}$
0.482±0.019±0.064		COFFMAN 88	MRK3	$e^+ e^- \rightarrow \pi^0 \pi^+ \pi^- \pi^0$

<sup>1</sup> Using  $B(\omega \rightarrow \pi^+ \pi^- \pi^0) = (89.1 \pm 0.7)\%$ .







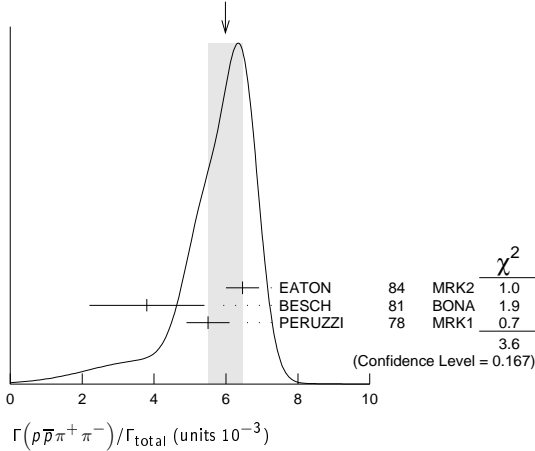
## Meson Particle Listings

 $J/\psi(1S)$ 

$\Gamma(p\bar{p}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{109}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>1.19 \pm 0.08</math> OUR AVERAGE</b>	Error	includes scale factor of 1.1.				
$1.33 \pm 0.02 \pm 0.11$	11k	ABLIKIM	09B	BES2	$e^+e^-$	
$1.13 \pm 0.09 \pm 0.09$	685	EATON	84	MRK2	$e^+e^-$	
$1.4 \pm 0.4$		BRANDELIK	79C	DASP	$e^+e^-$	
$1.00 \pm 0.15$	109	PERUZZI	78	MRK1	$e^+e^-$	

$\Gamma(p\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{110}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>6.0 \pm 0.5</math> OUR AVERAGE</b>	Error	includes scale factor of 1.3. See the ideogram below.				
$6.46 \pm 0.17 \pm 0.43$	1435	EATON	84	MRK2	$e^+e^-$	
$3.8 \pm 1.6$	48	BESCH	81	BONA	$e^+e^-$	
$5.5 \pm 0.6$	533	PERUZZI	78	MRK1	$e^+e^-$	

WEIGHTED AVERAGE  
 $6.0 \pm 0.5$  (Error scaled by 1.3)



$\Gamma(p\bar{p}\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{111}/\Gamma$	
Including $p\bar{p}\pi^+\pi^-\gamma$ and excluding $\omega, \eta, \eta'$						
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>2.3 \pm 0.9</math> OUR AVERAGE</b>	Error	includes scale factor of 1.9.				
$3.36 \pm 0.65 \pm 0.28$	364	EATON	84	MRK2	$e^+e^-$	
$1.6 \pm 0.6$	39	PERUZZI	78	MRK1	$e^+e^-$	

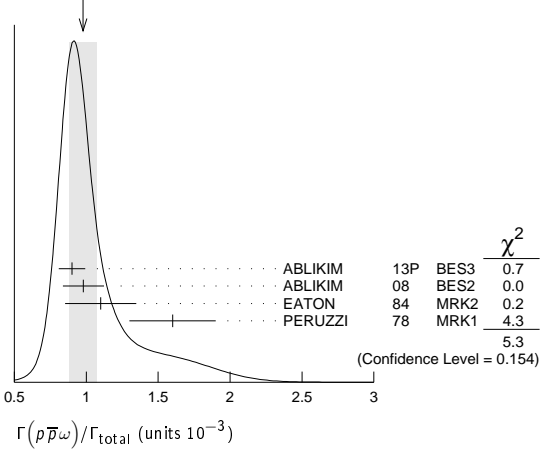
$\Gamma(p\bar{p}\eta)/\Gamma_{\text{total}}$					$\Gamma_{112}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>2.00 \pm 0.12</math> OUR AVERAGE</b>						
$1.91 \pm 0.02 \pm 0.17$	13k	<sup>1</sup> ABLIKIM	09	BES2	$e^+e^-$	
$2.03 \pm 0.13 \pm 0.15$	826	EATON	84	MRK2	$e^+e^-$	
$2.5 \pm 1.2$		BRANDELIK	79C	DASP	$e^+e^-$	
$2.3 \pm 0.4$	197	PERUZZI	78	MRK1	$e^+e^-$	

<sup>1</sup> From the combination of  $p\bar{p}\eta \rightarrow p\bar{p}\gamma\gamma$  and  $p\bar{p}\eta \rightarrow p\bar{p}\pi^+\pi^-\pi^0$  channels.

$\Gamma(p\bar{p}\rho)/\Gamma_{\text{total}}$					$\Gamma_{113}/\Gamma$	
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b><math>&lt; 0.31</math></b>	90	EATON	84	MRK2	$e^+e^- \rightarrow \text{hadrons}\gamma$	

$\Gamma(p\bar{p}\omega)/\Gamma_{\text{total}}$					$\Gamma_{114}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>0.98 \pm 0.10</math> OUR AVERAGE</b>	Error	includes scale factor of 1.3. See the ideogram below.				
$0.90 \pm 0.02 \pm 0.09$	2670	ABLIKIM	13P	BES3	$e^+e^-$	
$0.98 \pm 0.03 \pm 0.14$	2449	ABLIKIM	08	BES2	$e^+e^-$	
$1.10 \pm 0.17 \pm 0.18$	486	EATON	84	MRK2	$e^+e^-$	
$1.6 \pm 0.3$	77	PERUZZI	78	MRK1	$e^+e^-$	

WEIGHTED AVERAGE  
 $0.98 \pm 0.10$  (Error scaled by 1.3)



$\Gamma(p\bar{p}\eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{115}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>0.21 \pm 0.04</math> OUR AVERAGE</b>						
$0.200 \pm 0.023 \pm 0.028$	$265 \pm 31$	<sup>1</sup> ABLIKIM	09	BES2	$e^+e^-$	
$0.68 \pm 0.23 \pm 0.17$	19	EATON	84	MRK2	$e^+e^-$	
$1.8 \pm 0.6$	19	PERUZZI	78	MRK1	$e^+e^-$	

<sup>1</sup> From the combination of  $p\bar{p}\eta' \rightarrow p\bar{p}\pi^+\pi^-\eta$  and  $p\bar{p}\eta' \rightarrow p\bar{p}\gamma\rho^0$  channels.

$\Gamma(p\bar{p}a_0(980) \rightarrow p\bar{p}\pi^0\eta)/\Gamma_{\text{total}}$				$\Gamma_{116}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>6.8 \pm 1.2 \pm 1.3</math></b>	ABLIKIM	14N	BES3	$e^+e^- \rightarrow J/\psi$

$\Gamma(p\bar{p}\phi)/\Gamma_{\text{total}}$	$\Gamma_{117}/\Gamma$			
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.45 \pm 0.13 \pm 0.07</math></b>	FALVARD	88	DM2	$J/\psi \rightarrow \text{hadrons}$

$\Gamma(n\bar{n})/\Gamma_{\text{total}}$					$\Gamma_{118}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>2.09 \pm 0.16</math> OUR AVERAGE</b>						
$2.07 \pm 0.01 \pm 0.17$	36k	ABLIKIM	12C	BES3	$e^+e^-$	
$2.31 \pm 0.49$	79	BALDINI	98	FENI	$e^+e^-$	
$1.8 \pm 0.9$		BESCH	78	BONA	$e^+e^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$1.90 \pm 0.55$	40	ANTONELLI	93	SPEC	$e^+e^-$	

$\Gamma(n\bar{n}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{119}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>3.8 \pm 3.6</math></b>	5	BESCH	81	BONA	$e^+e^-$	

$\Gamma(\Sigma^+\Sigma^-)/\Gamma_{\text{total}}$					$\Gamma_{120}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>1.50 \pm 0.10 \pm 0.22</math></b>	399	ABLIKIM	08o	BES2	$e^+e^- \rightarrow J/\psi$	

$\Gamma(\Sigma^0\Sigma^0)/\Gamma_{\text{total}}$					$\Gamma_{121}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>1.29 \pm 0.09</math> OUR AVERAGE</b>						
$1.15 \pm 0.24 \pm 0.03$		<sup>1</sup> AUBERT	07BD	BABR	$10.6 e^+e^- \rightarrow \Sigma^0\Sigma^0\gamma$	
$1.33 \pm 0.04 \pm 0.11$	1779	ABLIKIM	06	BES2	$J/\psi \rightarrow \Sigma^0\bar{\Sigma}^0$	
$1.06 \pm 0.04 \pm 0.23$	$884 \pm 30$	PALLIN	87	DM2	$e^+e^- \rightarrow \Sigma^0\bar{\Sigma}^0$	
$1.58 \pm 0.16 \pm 0.25$	90	EATON	84	MRK2	$e^+e^- \rightarrow \Sigma^0\bar{\Sigma}^0$	
$1.3 \pm 0.4$	52	PERUZZI	78	MRK1	$e^+e^- \rightarrow \Sigma^0\bar{\Sigma}^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$2.4 \pm 2.6$	3	BESCH	81	BONA	$e^+e^- \rightarrow \Sigma^+\Sigma^-$	

<sup>1</sup> AUBERT 07BD reports  $[\Gamma(J/\psi(1S) \rightarrow \Sigma^0\bar{\Sigma}^0)/\Gamma_{\text{total}}] \times [\Gamma(J/\psi(1S) \rightarrow e^+e^-)] = (6.4 \pm 1.2 \pm 0.6) \times 10^{-3}$  keV which we divide by our best value  $\Gamma(J/\psi(1S) \rightarrow e^+e^-) = 5.55 \pm 0.14 \pm 0.02$  keV. Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2(\pi^+\pi^-)K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{122}/\Gamma$	
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>47 \pm 7</math> OUR AVERAGE</b>	Error	includes scale factor of 1.3.				
$49.8 \pm 4.2 \pm 3.4$	205	<sup>1</sup> AUBERT	06D	BABR	$10.6 e^+e^- \rightarrow \omega K^+ K^- 2(\pi^+\pi^-)\gamma$	
$31 \pm 13$	30	VANNUCCI	77	MRK1	$e^+e^-$	

<sup>1</sup> Using  $\Gamma(J/\psi \rightarrow e^+e^-) = 5.52 \pm 0.14 \pm 0.04$  keV.







$\Gamma(\gamma\rho\rho)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>4.5 ± 0.8 OUR AVERAGE</b>				
4.7 ± 0.3 ± 0.9		1 BALTRUSAIT...86B	MRK3	$J/\psi \rightarrow 4\pi\gamma$
3.75 ± 1.05 ± 1.20		2 BURKE	82	MRK2 $J/\psi \rightarrow 4\pi\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.09 90 3 BISELLO 89B  $J/\psi \rightarrow 4\pi\gamma$

<sup>1</sup> 4 $\pi$  mass less than 2.0 GeV.  
<sup>2</sup> 4 $\pi$  mass less than 2.0 GeV. We have multiplied 2 $\rho^0$  measurement by 3 to obtain 2 $\rho$ .  
<sup>3</sup> 4 $\pi$  mass in the range 2.0–25 GeV.

$\Gamma(\gamma\rho\omega)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.4</b>	90	ABLIKIM	08A	BES2 $e^+e^- \rightarrow J/\psi$

$\Gamma(\gamma\rho\phi)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;8.8</b>	90	ABLIKIM	08A	BES2 $e^+e^- \rightarrow J/\psi$

$\Gamma(\gamma\eta'(958))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>5.15 ± 0.16 OUR AVERAGE</b>				

Error includes scale factor of 1.2.

4.82 ± 0.23 ± 0.08 1 ABLIKIM 11 BES3  $J/\psi \rightarrow \eta'\gamma$

5.24 ± 0.12 ± 0.11 PEDLAR 09 CLE3  $J/\psi \rightarrow \eta'\gamma$

5.55 ± 0.44 35k ABLIKIM 06E BES2  $J/\psi \rightarrow \eta'\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.50 ± 0.14 ± 0.53 BOLTON 92B MRK3  $J/\psi \rightarrow \gamma\pi^+\pi^-\eta, \eta \rightarrow \gamma\gamma$

4.30 ± 0.31 ± 0.71 BOLTON 92B MRK3  $J/\psi \rightarrow \gamma\pi^+\pi^-\eta, \eta \rightarrow \pi^+\pi^-\pi^0$

4.04 ± 0.16 ± 0.85 622 AUGUSTIN 90 DM2  $J/\psi \rightarrow \gamma\eta\pi^+\pi^-$

4.39 ± 0.09 ± 0.66 2420 AUGUSTIN 90 DM2  $J/\psi \rightarrow \gamma\gamma\pi^+\pi^-$

4.1 ± 0.3 ± 0.6 BLOOM 83 CBAL  $e^+e^- \rightarrow 3\gamma + \text{hadrons}$

2.9 ± 1.1 6 BRANDELIK 79c DASP  $e^+e^- \rightarrow 3\gamma$

2.4 ± 0.7 57 BARTEL 76 CNTR  $e^+e^- \rightarrow 2\gamma\rho$

<sup>1</sup> ABLIKIM 11 reports  $(4.84 \pm 0.03 \pm 0.24) \times 10^{-3}$  from a measurement of  $[\Gamma(J/\psi(1S) \rightarrow \gamma\eta'(958))/\Gamma_{\text{total}}] / [B(\eta'(958) \rightarrow \pi^+\pi^-\eta)] / [B(\eta \rightarrow 2\gamma)]$  assuming  $B(\eta'(958) \rightarrow \pi^+\pi^-\eta) = (43.2 \pm 0.7) \times 10^{-2}$ ,  $B(\eta \rightarrow 2\gamma) = (39.31 \pm 0.20) \times 10^{-2}$ , which we rescale to our best values  $B(\eta'(958) \rightarrow \pi^+\pi^-\eta) = (42.9 \pm 0.7) \times 10^{-2}$ ,  $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

$\Gamma(\gamma 2\pi^+ 2\pi^-)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.8 ± 0.5 OUR AVERAGE</b>				

Error includes scale factor of 1.9. See the ideogram below.

4.32 ± 0.14 ± 0.73 1 BISELLO 89B DM2  $J/\psi \rightarrow 4\pi\gamma$

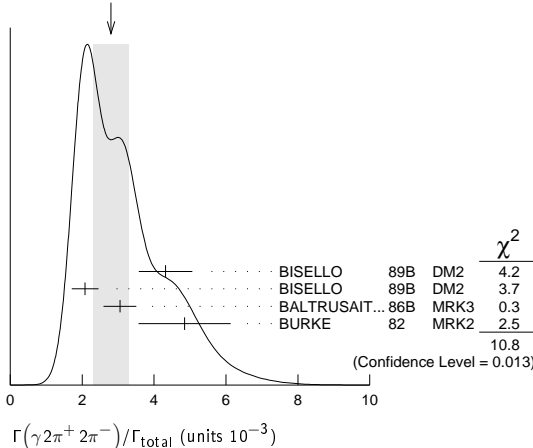
2.08 ± 0.13 ± 0.35 2 BISELLO 89B DM2  $J/\psi \rightarrow 4\pi\gamma$

3.05 ± 0.08 ± 0.45 2 BALTRUSAIT...86B MRK3  $J/\psi \rightarrow 4\pi\gamma$

4.85 ± 0.45 ± 1.20 3 BURKE 82 MRK2  $e^+e^-$

<sup>1</sup> 4 $\pi$  mass less than 3.0 GeV.  
<sup>2</sup> 4 $\pi$  mass less than 2.0 GeV.  
<sup>3</sup> 4 $\pi$  mass less than 2.5 GeV.

WEIGHTED AVERAGE  
2.8 ± 0.5 (Error scaled by 1.9)



$\Gamma(\gamma f_2(1270) f_2(1270))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>9.5 ± 0.7 ± 1.6</b>	646 ± 45	ABLIKIM	04M	BES $J/\psi \rightarrow \gamma 2\pi^+ 2\pi^-$

$\Gamma_{160}/\Gamma$

$\Gamma(\gamma f_2(1270) f_2(1270) (\text{non resonant}))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>8.2 ± 0.8 ± 1.7</b>				

<sup>1</sup> Subtracting contribution from intermediate  $\eta_c(1S)$  decays.

$\Gamma(\gamma K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.1 ± 0.1 ± 0.6</b>	1516	BAI	00B	BES $J/\psi \rightarrow \gamma K^+ K^0 \pi^+ \pi^-$

$\Gamma(\gamma f_4(2050))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>2.7 ± 0.5 ± 0.5</b>				

<sup>1</sup> Assuming branching fraction  $f_4(2050) \rightarrow \pi\pi/\text{total} = 0.167$ .

$\Gamma(\gamma\omega\omega)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.61 ± 0.33 OUR AVERAGE</b>				

6.0 ± 4.8 ± 1.8 ABLIKIM 08A BES2  $J/\psi \rightarrow \gamma\omega\pi^+\pi^-$

1.41 ± 0.2 ± 0.42 120 ± 17 BISELLO 87 SPEC  $e^+e^-$ , hadrons  $\gamma$

1.76 ± 0.09 ± 0.45 BALTRUSAIT...85C MRK3  $e^+e^- \rightarrow \text{hadrons}\gamma$

$\Gamma(\gamma\eta(1405/1475) \rightarrow \gamma\rho^0\rho^0)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.7 ± 0.4 OUR AVERAGE</b>				

Error includes scale factor of 1.3.

2.1 ± 0.4 BUGG 95 MRK3  $J/\psi \rightarrow \gamma\pi^+\pi^-\pi^+\pi^-$

1.36 ± 0.38 1,2 BISELLO 89B DM2  $J/\psi \rightarrow 4\pi\gamma$

<sup>1</sup> Estimated by us from various fits.  
<sup>2</sup> Includes unknown branching fraction to  $\rho^0\rho^0$ .

$\Gamma(\gamma f_2(1270))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT
<b>1.64 ± 0.12 OUR AVERAGE</b>				

Error includes scale factor of 1.3. See the ideogram below.

2.07 ± 0.16 ± 0.02 2.4k 1,2 DOBBS 15  $J/\psi \rightarrow \gamma\pi\pi$

1.63 ± 0.26 ± 0.02 3 ABLIKIM 06v BES2  $e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^+\pi^-$

1.42 ± 0.21 ± 0.01 4 ABLIKIM 06v BES2  $e^+e^- \rightarrow J/\psi \rightarrow \gamma\pi^0\pi^0$

1.33 ± 0.05 ± 0.20 5 AUGUSTIN 87 DM2  $J/\psi \rightarrow \gamma\pi^+\pi^-$

1.36 ± 0.09 ± 0.23 5 BALTRUSAIT...87 MRK3  $J/\psi \rightarrow \gamma\pi^+\pi^-$

1.48 ± 0.25 ± 0.30 178 EDWARDS 82B CBAL  $e^+e^- \rightarrow 2\pi^0\gamma$

2.0 ± 0.7 35 ALEXANDER 78 PLUT  $e^+e^-$

1.2 ± 0.6 30 6 BRANDELIK 78B DASP  $e^+e^- \rightarrow \pi^+\pi^-\gamma$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.

<sup>2</sup> DOBBS 15 reports  $[\Gamma(J/\psi(1S) \rightarrow \gamma f_2(1270))/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] = (1.744 \pm 0.052 \pm 0.122) \times 10^{-3}$  which we divide by our best value  $B(f_2(1270) \rightarrow \pi\pi) = (84.2^{+2.9}_{-0.9}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

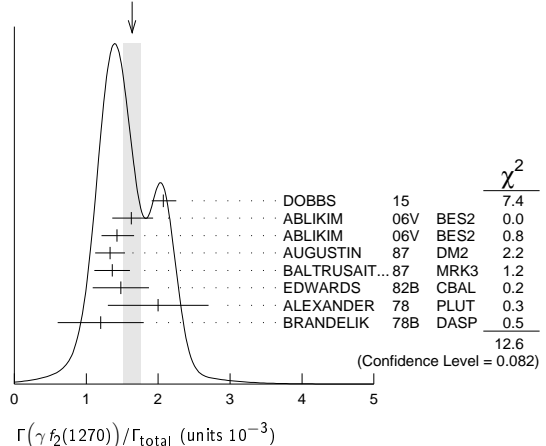
<sup>3</sup> ABLIKIM 06v reports  $[\Gamma(J/\psi(1S) \rightarrow \gamma f_2(1270))/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] = (1.371 \pm 0.010 \pm 0.222) \times 10^{-3}$  which we divide by our best value  $B(f_2(1270) \rightarrow \pi\pi) = (84.2^{+2.9}_{-0.9}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> ABLIKIM 06v reports  $[\Gamma(J/\psi(1S) \rightarrow \gamma f_2(1270))/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] = (1.200 \pm 0.027 \pm 0.174) \times 10^{-3}$  which we divide by our best value  $B(f_2(1270) \rightarrow \pi\pi) = (84.2^{+2.9}_{-0.9}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>5</sup> Estimated using  $B(f_2(1270) \rightarrow \pi\pi) = 0.843 \pm 0.012$ . The errors do not contain the uncertainty in the  $f_2(1270)$  decay.

<sup>6</sup> Restated by us to take account of spread of E1, M2, E3 transitions.

WEIGHTED AVERAGE  
1.64 ± 0.12 (Error scaled by 1.3)











# Meson Particle Listings

## $J/\psi(1S)$ , Branching Ratios of $\psi$ 's and $\chi$ 's, $\chi_{c0}(1P)$

FRANKLIN	83	PRL 51 963	M.E.B. Franklin <i>et al.</i>	(LBL, SLAC)
BURKE	82	PRL 49 632	D.L. Burke <i>et al.</i>	(LBL, SLAC)
EDWARDS	82B	PR D25 3065	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
EDWARDS	82D	PRL 48 458	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
Also		ARNS 33 143	E.D. Bloom, C. Peck	(SLAC, CIT)
EDWARDS	82E	PRL 49 259	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)
LEMOIGNE	82	PL 113B 509	Y. Lemoigne <i>et al.</i>	(SACL, LOIC, SHMP+)
BESCH	81	ZPHY C8 1	H.J. Besch <i>et al.</i>	(BONN, DESY, MANZ)
GIDAL	81	PL 107B 153	G. Gidal <i>et al.</i>	(SLAC, LBL)
PATRIDGE	80	PRL 44 712	R. Patridge <i>et al.</i>	(CIT, HARV, PRIN+)
SCHARRE	80	PL 97B 329	D.L. Scharre <i>et al.</i>	(SLAC, LBL)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
Also		SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)
BRANDELIC	79C	ZPHY C1 233	R. Brandelik <i>et al.</i>	(DASP Collab.)
ALEXANDER	78	PL 72B 493	G. Alexander <i>et al.</i>	(DESY, HAMB, SIEG+)
BESCH	78	PL 78B 347	H.J. Besch <i>et al.</i>	(BONN, DESY, MANZ)
BRANDELIC	78B	PL 74B 292	R. Brandelik <i>et al.</i>	(DASP Collab.)
PERUZZI	78	PR D17 2901	I. Peruzzi <i>et al.</i>	(SLAC, LBL)
BARTEL	77	PL 66B 489	W. Bartel <i>et al.</i>	(DESY, HEIDP)
BURMESTER	77D	PL 72B 135	J. Burmester <i>et al.</i>	(DESY, HAMB, SIEG+)
FELDMAN	77	PRPL 33C 285	G.J. Feldman, M.L. Perl	(LBL, SLAC)
VANNUCCI	77	PR D15 1814	F. Vannucci <i>et al.</i>	(SLAC, LBL)
BARTEL	76	PL 64B 483	W. Bartel <i>et al.</i>	(DESY, HEIDP)
BRAUNSCH...	76	PL 63B 487	W. Braunschweig <i>et al.</i>	(DASP Collab.)
JEAN-MARIE	76	PRL 36 291	B. Jean-Marie <i>et al.</i>	(SLAC, LBL) IG
BALDINI...	75	PL 58B 471	R. Baldini-Celio <i>et al.</i>	(FRAS, ROMA)
BOYARSKI	75	PRL 34 1357	A.M. Boyarski <i>et al.</i>	(SLAC, LBL) JPC
DASP	75	PL 56B 491	W. Braunschweig <i>et al.</i>	(DASP Collab.)
ESPOSITO	75B	LNC 14 73	B. Esposito <i>et al.</i>	(FRAS, NAPL, PADO+)
FORD	75	PRL 34 604	R.L. Ford <i>et al.</i>	(SLAC, PENN)

### BRANCHING RATIOS OF $\psi(2S)$ AND $\chi_{c0,1,2}$

Updated August 2015 by J.J. Hernández-Rey (IFIC, Valencia), S. Navas (University of Granada), and C. Patrignani (INFN, Bologna Univ.)

Since 2002, the treatment of the branching ratios of the  $\psi(2S)$  and  $\chi_{c0,1,2}$  has undergone an important restructuring.

When measuring a branching ratio experimentally, it is not always possible to normalize the number of events observed in the corresponding decay mode to the total number of particles produced. Therefore, the experimenters sometimes report the number of observed decays with respect to another decay mode of the same or another particle in the relevant decay chain. This is actually equivalent to measuring combinations of branching fractions of several decay modes.

To extract the branching ratio of a given decay mode, the collaborations use some previously reported measurements of the required branching ratios. However, the values are frequently taken from the *Review of Particle Physics* (RPP), which in turn uses the branching ratio reported by the experiment in the following edition, giving rise either to correlations or to plain vicious circles Ref. 1, Ref. 2 as discussed in more detail in earlier editions of this mini-review.

The way to avoid these dependencies and correlations is to extract the branching ratios through a fit that uses the truly measured combinations of branching fractions and partial widths. This fit, in fact, should involve decays from the four concerned particles,  $\psi(2S)$ ,  $\chi_{c0}$ ,  $\chi_{c1}$ , and  $\chi_{c2}$ , and occasionally some combinations of branching ratios of more than one of them. This is what is done since the 2002 edition [3].

The PDG policy is to quote the results of the collaborations in a manner as close as possible to what appears in their original publications. However, in order to avoid the problems mentioned above, we had in some cases to work out the values originally measured, using the number of events and detection efficiencies given by the collaborations, or rescaling back the published results. The information was sometimes spread over several articles, and some articles referred to papers still unpublished, which in turn contained the relevant numbers in footnotes.

Even though the experimental collaborations are entitled to extract whatever branching ratios they consider appropriate by

using other published results, we would like to encourage them to also quote explicitly in their articles the actual quantities measured, so that they can be used directly in averages and fits of different experimental determinations.

To inform the reader how we computed some of the values used in this edition of RPP, we use footnotes to indicate the branching ratios actually given by the experiments and the quantities they use to derive them from the true combination of branching ratios actually measured.

None of the branching ratios of the  $\chi_{c0,1,2}$  are measured independently of the  $\psi(2S)$  radiative decays. We tried to identify those branching ratios which can be correlated in a non-trivial way, and although we cannot preclude the existence of other cases, we are confident that the most relevant correlations have already been removed. Nevertheless, correlations in the errors of different quantities measured by the same experiment have not been taken into account.

### FIT INFORMATION

This is an overall fit to 4 total widths, 1 partial width, 25 combinations of partial widths, 10 branching ratios, and 80 combinations of branching ratios. Of the latter 57 involve decays of more than one particle.

The overall fit uses 239 measurements to determine 49 parameters and has a  $\chi^2$  of 341.3 for 190 degrees of freedom.

The relatively high  $\chi^2$  of the fit, 1.8 per d.o.f., can be traced back to a few specific discrepancies in the data. No scaling factors to fit uncertainties have been applied.

In the listing we provide the inter-particle correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the corresponding parameter  $x_i$ .

### References

1. Y.F. Gu and X.H. Li, Phys. Lett. **B449**, 361 (1999).
2. C. Patrignani, Phys. Rev. **D64**, 034017 (2001).
3. Particle Data Group, K.Hagiwara *et al.*, Phys. Rev. **D68**, 010001 (2002).

$\chi_{c0}(1P)$

$I^G(J^{PC}) = 0^+(0^{++})$

$\chi_{c0}(1P)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3414.75 ± 0.31 OUR AVERAGE</b>				
3414.2 ± 0.5 ± 2.3	5.4k	UEHARA	08 BELL	$\gamma\gamma \rightarrow \chi_{c0} \rightarrow \text{hadrons}$
3406 ± 7 ± 6	230	<sup>1</sup> ABE	07 BELL	$e^+e^- \rightarrow J/\psi(c\bar{c})$
3414.21 ± 0.39 ± 0.27		ABLIKIM	05G BES2	$\psi(2S) \rightarrow \gamma\chi_{c0}$
3414.7 ± 0.7 ± 0.2		<sup>2</sup> ANDREOTTI	03 E835	$\bar{p}p \rightarrow \chi_{c0} \rightarrow \pi^0\pi^0$
3415.5 ± 0.4 ± 0.4	392	<sup>3</sup> BAGNASCO	02 E835	$\bar{p}p \rightarrow \chi_{c0} \rightarrow J/\psi\gamma$
3417.4 ± 1.8 ± 1.9		<sup>2</sup> AMBROGIANI	99B E835	$\bar{p}p \rightarrow e^+e^-\gamma$
3414.1 ± 0.6 ± 0.8		BAI	99B BES	$\psi(2S) \rightarrow \gamma X$
3417.8 ± 0.4 ± 4		<sup>2</sup> GAISER	86 CBAL	$\psi(2S) \rightarrow \gamma X$
3416 ± 3 ± 4		<sup>4</sup> TANENBAUM	78 MRK1	$e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3414.6 ± 1.1	266	UEHARA	13 BELL	$\gamma\gamma \rightarrow K_S^0 K_S^0$
3416.5 ± 3.0		EISENSTEIN	01 CLE2	$e^+e^- \rightarrow e^+e^-\chi_{c0}$
3422 ± 10		<sup>4</sup> BARTEL	78B CNTR	$e^+e^- \rightarrow J/\psi 2\gamma$
3415 ± 9		<sup>4</sup> BIDDICK	77 CNTR	$e^+e^- \rightarrow \gamma X$

<sup>1</sup> From a fit of the  $J/\psi$  recoil mass spectrum. Supersedes ABE,K 02 and ABE 04g.  
<sup>2</sup> Using mass of  $\psi(2S) = 3686.0$  MeV.







$\Gamma(\pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID
<b><math>17.5 \pm 1.4</math> OUR FIT</b>	

$\Gamma(K^+\bar{K}^*(892)^0\pi^- + \text{c.c.})/\Gamma(\pi^+\pi^-K^+K^-)$	$\Gamma_{30}/\Gamma_8$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.41 \pm 0.09</math> OUR FIT</b>	
<b><math>0.41 \pm 0.10</math></b>	TANENBAUM 78 MRK1 $\psi(2S) \rightarrow \gamma\chi_{c0}$

$\Gamma(K_0^*(1430)^0\bar{K}_0^*(1430)^0 \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
VALUE (units $10^{-4}$ ) EVTS	DOCUMENT ID TECN COMMENT
<b><math>9.6 \pm 2.8 \pm 0.3</math></b>	83 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $(10.44 \pm 2.37^{+3.05}_{-1.90}) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow K_0^*(1430)^0\bar{K}_0^*(1430)^0 \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K_0^*(1430)^0\bar{K}_2^*(1430)^0 + \text{c.c.} \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
VALUE (units $10^{-4}$ ) EVTS	DOCUMENT ID TECN COMMENT
<b><math>7.8 \pm 2.9 \pm 0.2</math></b>	62 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $(8.49 \pm 1.66^{+1.32}_{-1.99}) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow K_0^*(1430)^0\bar{K}_2^*(1430)^0 + \text{c.c.} \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K_1(1270)^+K^-\text{ + c.c.} \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$
VALUE (units $10^{-3}$ ) EVTS	DOCUMENT ID TECN COMMENT
<b><math>6.1 \pm 1.9 \pm 0.2</math></b>	68 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $(6.66 \pm 1.31^{+1.60}_{-1.51}) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow K_1(1270)^+K^- + \text{c.c.} \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. The measurement assumes  $\text{B}(K_1(1270) \rightarrow K\rho(770)) = 42 \pm 6\%$ .

$\Gamma(K_1(1400)^+K^-\text{ + c.c.} \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$
VALUE (units $10^{-3}$ ) CL% DOCUMENT ID TECN COMMENT	
<b><math>&lt;2.6</math></b>	90 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $< 2.85 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow K_1(1400)^+K^- + \text{c.c.} \rightarrow \pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . The measurement assumes  $\text{B}(K_1(1400) \rightarrow K^*(892)\pi) = 94 \pm 6\%$ .

$\Gamma(f_0(980)f_0(980))/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$
VALUE (units $10^{-5}$ ) EVTS	DOCUMENT ID TECN COMMENT
<b><math>15.9 \pm 10.2 \pm 0.4</math></b>	28 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(980)f_0(980))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$   $= (1.59 \pm 0.50^{+0.89}_{-0.72}) \times 10^{-5}$  which we divide by our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. One of the  $f_0(980)$  mesons is identified via decay to  $\pi^+\pi^-$  while the other via  $K^+K^-$  decay.

$\Gamma(f_0(980)f_0(2200))/\Gamma_{\text{total}}$	$\Gamma_{14}/\Gamma$
VALUE (units $10^{-4}$ ) EVTS	DOCUMENT ID TECN COMMENT
<b><math>7.8 \pm 2.9 \pm 0.2</math></b>	77 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $(8.42 \pm 1.42^{+1.65}_{-2.29}) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(980)f_0(2200))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. The  $f_0$  mesons are identified via  $f_0(980) \rightarrow \pi^+\pi^-$  and  $f_0(2200) \rightarrow K^+K^-$  decays.

$\Gamma(f_0(1370)f_0(1370))/\Gamma_{\text{total}}$	$\Gamma_{15}/\Gamma$
VALUE (units $10^{-4}$ ) CL% DOCUMENT ID TECN COMMENT	
<b><math>&lt;2.7</math></b>	90 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $< 2.9 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(1370)f_0(1370))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . One of the  $f_0(1370)$  mesons is identified via decay to  $\pi^+\pi^-$  while the other via  $K^+K^-$  decay. Both branching fractions for these  $f_0$  decays are implicitly included in the quoted result.

$\Gamma(f_0(1370)f_0(1500))/\Gamma_{\text{total}}$	$\Gamma_{16}/\Gamma$
VALUE (units $10^{-4}$ ) CL% DOCUMENT ID TECN COMMENT	
<b><math>&lt;1.7</math></b>	90 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $< 1.8 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(1370)f_0(1500))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . The  $f_0$  mesons are identified via  $f_0(1370) \rightarrow \pi^+\pi^-$  and  $f_0(1500) \rightarrow K^+K^-$  decays. Both branching fractions for these  $f_0$  decays are implicitly included in the quoted result.

$\Gamma(f_0(1370)f_0(1710))/\Gamma_{\text{total}}$	$\Gamma_{17}/\Gamma$
VALUE (units $10^{-4}$ ) EVTS DOCUMENT ID TECN COMMENT	
<b><math>6.6 \pm 3.5 \pm 0.2</math></b>	61 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $(7.12 \pm 1.85^{+3.28}_{-1.68}) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(1370)f_0(1710))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. The  $f_0$  mesons are identified via  $f_0(1370) \rightarrow \pi^+\pi^-$  and  $f_0(1710) \rightarrow K^+K^-$  decays. Both branching fractions for these  $f_0$  decays are implicitly included in the quoted result.

$\Gamma(f_0(1500)f_0(1370))/\Gamma_{\text{total}}$	$\Gamma_{18}/\Gamma$
VALUE (units $10^{-4}$ ) CL% DOCUMENT ID TECN COMMENT	
<b><math>&lt;1.3</math></b>	90 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $< 1.4 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(1500)f_0(1370))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . The  $f_0$  mesons are identified via  $f_0(1500) \rightarrow \pi^+\pi^-$  and  $f_0(1370) \rightarrow K^+K^-$  decays. Both branching fractions for these  $f_0$  decays are implicitly included in the quoted result.

$\Gamma(f_0(1500)f_0(1500))/\Gamma_{\text{total}}$	$\Gamma_{19}/\Gamma$
VALUE (units $10^{-4}$ ) CL% DOCUMENT ID TECN COMMENT	
<b><math>&lt;0.5</math></b>	90 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $< 0.55 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(1500)f_0(1500))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . One of the  $f_0(1500)$  is identified via decay to  $\pi^+\pi^-$  while the other via  $K^+K^-$  decay. Both branching fractions for these  $f_0$  decays are implicitly included in the quoted result.

$\Gamma(f_0(1500)f_0(1710))/\Gamma_{\text{total}}$	$\Gamma_{20}/\Gamma$
VALUE (units $10^{-4}$ ) CL% DOCUMENT ID TECN COMMENT	
<b><math>&lt;0.7</math></b>	90 <sup>1</sup> ABLIKIM 05Q BES2 $\psi(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-$

<sup>1</sup> ABLIKIM 05Q reports  $< 0.73 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow f_0(1500)f_0(1710))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . The  $f_0$  mesons are identified via  $f_0(1500) \rightarrow \pi^+\pi^-$  and  $f_0(1710) \rightarrow K^+K^-$  decays. Both branching fractions for these  $f_0$  decays are implicitly included in the quoted result.

$\Gamma(K^+K^-\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{21}/\Gamma$
VALUE (units $10^{-3}$ ) EVTS DOCUMENT ID TECN COMMENT	
<b><math>8.61 \pm 0.13 \pm 0.94</math></b>	9.0k <sup>1</sup> ABLIKIM 13B BES3 $e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c0}$

<sup>1</sup> Using  $1.06 \times 10^8$   $\psi(2S)$  mesons and  $\text{B}(\psi(2S) \rightarrow \chi_{c0}\gamma) = (9.68 \pm 0.31)\%$ .

$\Gamma(K_S^0K^{\pm}\pi^{\mp}\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{22}/\Gamma$
VALUE (units $10^{-3}$ ) EVTS DOCUMENT ID TECN COMMENT	
<b><math>4.22 \pm 0.10 \pm 0.43</math></b>	2.7k <sup>1</sup> ABLIKIM 13B BES3 $e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c0}$

<sup>1</sup> Using  $1.06 \times 10^8$   $\psi(2S)$  mesons and  $\text{B}(\psi(2S) \rightarrow \chi_{c0}\gamma) = (9.68 \pm 0.31)\%$ .

$\Gamma(K^+K^-\pi^0\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{23}/\Gamma$
VALUE (%) EVTS DOCUMENT ID TECN COMMENT	
<b><math>0.54 \pm 0.09 \pm 0.01</math></b>	213.5 <sup>1</sup> HE 08B CLEO $e^+e^- \rightarrow \gamma h^+h^-h^0h^0$

<sup>1</sup> HE 08B reports  $0.59 \pm 0.05 \pm 0.08 \pm 0.03\%$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow K^+K^-\pi^0\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{24}/\Gamma$
VALUE (%) EVTS DOCUMENT ID TECN COMMENT	
<b><math>2.44 \pm 0.32 \pm 0.07</math></b>	401.7 <sup>1</sup> HE 08B CLEO $e^+e^- \rightarrow \gamma h^+h^-h^0h^0$

<sup>1</sup> HE 08B reports  $2.64 \pm 0.15 \pm 0.31 \pm 0.14\%$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.



$\Gamma(K_S^0 K_S^0)/\Gamma(K^+ K^-)$   $\Gamma_{43}/\Gamma_{42}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.52±0.04 OUR FIT</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.49±0.07±0.08	<sup>1,2</sup> CHEN	07B BELL	$e^+ e^- \rightarrow e^+ e^- \chi_{c0}$

<sup>1</sup> Using  $\Gamma(K^+ K^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  from NAKAZAWA 05.<sup>2</sup> Not independent from other measurements. $\Gamma(\pi^+ \pi^- \eta)/\Gamma_{\text{total}}$   $\Gamma_{44}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.19</b>	90	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.0	90	<sup>2</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c0}$

<sup>1</sup> ATHAR 07 reports  $< 0.21 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \pi^+ \pi^- \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .<sup>2</sup> ABLIKIM 06R reports  $< 1.1 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \pi^+ \pi^- \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ . $\Gamma(\pi^+ \pi^- \eta')/\Gamma_{\text{total}}$   $\Gamma_{45}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.35</b>	90	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$
<sup>1</sup> ATHAR 07 reports $< 0.38 \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow \pi^+ \pi^- \eta')/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .				

 $\Gamma(\bar{K}^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{46}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.09</b>	90	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.6	90	<sup>2,3</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c0}$
<0.7	90	<sup>3,4</sup> BAI	99B BES	$\psi(2S) \rightarrow \gamma \chi_{c0}$

<sup>1</sup> ATHAR 07 reports  $< 0.10 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .<sup>2</sup> ABLIKIM 06R reports  $< 0.70 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .<sup>3</sup> We have multiplied the  $K_S^0 K^+ \pi^-$  measurement by a factor of 2 to convert to  $K^0 K^+ \pi^-$ .<sup>4</sup> Rescaled by us using  $B(\psi(2S) \rightarrow \gamma \chi_{c0}) = (9.4 \pm 0.4)\%$  and  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (32.6 \pm 0.5)\%$ . $\Gamma(K^+ K^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{47}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.06</b>	90	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$
<sup>1</sup> ATHAR 07 reports $< 0.06 \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .				

 $\Gamma(K^+ K^- \eta)/\Gamma_{\text{total}}$   $\Gamma_{48}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.22</b>	90	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$
<sup>1</sup> ATHAR 07 reports $< 0.24 \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow K^+ K^- \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .				

 $\Gamma(K^+ K^- K_S^0 K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{49}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.38±0.46±0.04</b>	16.8±4.8	<sup>1</sup> ABLIKIM	05o BES2	$\psi(2S) \rightarrow \gamma \chi_{c0}$

<sup>1</sup> ABLIKIM 05o reports  $[\Gamma(\chi_{c0}(1P) \rightarrow K^+ K^- K_S^0 K_S^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  =  $(0.138 \pm 0.039 \pm 0.025) \times 10^{-3}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. $\Gamma(K^+ K^- K^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{50}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID
<b>2.75±0.28 OUR FIT</b>	

 $\Gamma(K^+ K^- \phi)/\Gamma_{\text{total}}$   $\Gamma_{51}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.95±0.24±0.03</b>	38	<sup>1</sup> ABLIKIM	06T BES2	$\psi(2S) \rightarrow \gamma 2K^+ 2K^-$
<sup>1</sup> ABLIKIM 06T reports $(1.03 \pm 0.22 \pm 0.15) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow K^+ K^- \phi)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

 $\Gamma(\bar{K}^0 K^+ \pi^- \phi + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{52}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>3.68±0.30±0.50</b>	ABLIKIM	15M BES3	$\psi(2S) \rightarrow \gamma \chi_{c0}$

 $\Gamma(K^+ K^- \pi^0 \phi)/\Gamma_{\text{total}}$   $\Gamma_{53}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.90±0.14±0.32</b>	ABLIKIM	15M BES3	$\psi(2S) \rightarrow \gamma \chi_{c0}$

 $\Gamma(\phi \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{54}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.18±0.07±0.13</b>	538	<sup>1</sup> ABLIKIM	13B BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c0}$
<sup>1</sup> Using $1.06 \times 10^8$ $\psi(2S)$ mesons and $B(\psi(2S) \rightarrow \chi_{c0} \gamma) = (9.68 \pm 0.31)\%$ .				

 $\Gamma(\phi \phi)/\Gamma_{\text{total}}$   $\Gamma_{55}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID
<b>0.77±0.07 OUR FIT</b>	

 $\Gamma(p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{56}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>2.25±0.09 OUR FIT</b>	

 $\Gamma(p\bar{p} \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{57}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.68±0.07 OUR AVERAGE</b>	Error includes scale factor of 1.3.		
0.72±0.06±0.02	<sup>1</sup> ONYISI	10 CLE3	$\psi(2S) \rightarrow \gamma p \bar{p} X$
0.54±0.11±0.01	<sup>2</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ONYISI 10 reports  $(7.76 \pm 0.37 \pm 0.51 \pm 0.39) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p} \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>2</sup> ATHAR 07 reports  $(0.59 \pm 0.10 \pm 0.08) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p} \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. $\Gamma(p\bar{p} \eta)/\Gamma_{\text{total}}$   $\Gamma_{58}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.35±0.04 OUR AVERAGE</b>			
0.34±0.04±0.01	<sup>1</sup> ONYISI	10 CLE3	$\psi(2S) \rightarrow \gamma p \bar{p} X$
0.36±0.11±0.01	<sup>2</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ONYISI 10 reports  $(3.73 \pm 0.38 \pm 0.28 \pm 0.19) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p} \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>2</sup> ATHAR 07 reports  $(0.39 \pm 0.11 \pm 0.04) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p} \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. $\Gamma(p\bar{p} \omega)/\Gamma_{\text{total}}$   $\Gamma_{59}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>0.51±0.05±0.01</b>	<sup>1</sup> ONYISI	10 CLE3	$\psi(2S) \rightarrow \gamma p \bar{p} X$

<sup>1</sup> ONYISI 10 reports  $(5.57 \pm 0.48 \pm 0.42 \pm 0.14) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p} \omega)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. $\Gamma(p\bar{p} \phi)/\Gamma_{\text{total}}$   $\Gamma_{60}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.9±1.4±0.2</b>	42±8	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p \bar{p} K^+ K^-$

<sup>1</sup> ABLIKIM 11F reports  $(6.12 \pm 1.18 \pm 0.86) \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p} \phi)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

## $\chi_{c0}(1P)$

$(p\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{61}/\Gamma$		
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.1 <math>\pm</math> 0.7 OUR EVALUATION</b>	Error includes scale factor of 1.4. Treating systematic error as correlated.		
<b>2.1 <math>\pm</math> 1.0 OUR AVERAGE</b>	Error includes scale factor of 2.0.		
1.57 $\pm$ 0.21 $\pm$ 0.53	<sup>1</sup> BAI	99B	BES $\psi(2S) \rightarrow \gamma\chi_{c0}$
4.20 $\pm$ 1.15 $\pm$ 0.18	<sup>1</sup> TANENBAUM	78	MRK1 $\psi(2S) \rightarrow \gamma\chi_{c0}$
<sup>1</sup> Rescaled by us using $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.4 \pm 0.4)\%$ and $B(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = (32.6 \pm 0.5)\%$ .			

$(p\overline{p}\pi^0\pi^0)/\Gamma_{\text{total}}$	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{62}/\Gamma$
VALUE (%)					
<b><math>0.102 \pm 0.027 \pm 0.003</math></b>	39.5	<sup>1</sup> HE	08B	CLEO $e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	
<sup>1</sup> HE 08B reports $0.11 \pm 0.02 \pm 0.02 \pm 0.01\%$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow p\overline{p}\pi^0\pi^0)/\Gamma_{\text{total}} \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(p\overline{p}K^+K^-(\text{non-resonant}))/\Gamma_{\text{total}}$				$\Gamma_{63}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.19 \pm 0.26 \pm 0.03</math></b>	$48 \pm 8$	<sup>1</sup> ABLIKIM	11F	BES3 $\psi(2S) \rightarrow \gamma p\overline{p}K^+K^-$
<sup>1</sup> ABLIKIM 11F reports $(1.24 \pm 0.20 \pm 0.18) \times 10^{-4}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow p\overline{p}K^+K^-(\text{non-resonant}))/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(p\overline{p}K_S^0K_S^0)/\Gamma_{\text{total}}$				$\Gamma_{64}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;8.8</b>	90	<sup>1</sup> ABLIKIM	06D BES2	$\psi(2S) \rightarrow \chi_{c0} \gamma$
<sup>1</sup> Using $B(\psi(2S) \rightarrow \chi_{c0} \gamma) = (9.2 \pm 0.5)\%$				

$(p\overline{p}\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{65}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>12.4 \pm 1.1</math> OUR AVERAGE</b>				
$12.6 \pm 1.1 \pm 0.3$	5150	<sup>1</sup> ABLIKIM	12j	BES3 $\psi(2S) \rightarrow \gamma p\overline{p}\pi^-$
$11.0 \pm 3.0 \pm 0.3$		<sup>2</sup> ABLIKIM	06i	BES2 $\psi(2S) \rightarrow \gamma p\pi^- X$
<sup>1</sup> ABLIKIM 12j reports $[\Gamma(\chi_{c0}(1P) \rightarrow p\overline{p}\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ = $(1.26 \pm 0.02 \pm 0.11) \times 10^{-4}$ which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> ABLIKIM 06i reports $[\Gamma(\chi_{c0}(1P) \rightarrow p\overline{p}\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ = $(1.10 \pm 0.24 \pm 0.18) \times 10^{-4}$ which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(\overline{p}n\pi^+)/\Gamma_{\text{total}}$	$\Gamma_{66}/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>13.4 \pm 1.1 \pm 0.4</math></b>	5808	<sup>1</sup> ABLIKIM	12j	BES3 $\psi(2S) \rightarrow \gamma \overline{p}n\pi^+$
<sup>1</sup> ABLIKIM 12j reports $[\Gamma(\chi_{c0}(1P) \rightarrow \overline{p}n\pi^+)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ = $(1.34 \pm 0.03 \pm 0.11) \times 10^{-4}$ which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))$ = $(9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$(\rho\pi^-\pi^+\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{67}/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>22.9 \pm 2.0 \pm 0.6</math></b>	2480	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\rho\pi^-\pi^+\pi^0$

<sup>1</sup>ABLIKIM 12i reports  $[\Gamma(\chi_{c0}(1P) \rightarrow \rho\pi^-\pi^+\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))] = (2.29 \pm 0.08 \pm 0.18) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\overline{p}n\pi^+\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{68}/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>21.6 \pm 1.7 \pm 0.6</math></b>	2757	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma \overline{p}n\pi^+\pi^0$
<sup>1</sup> ABLIKIM 12i reports $[\Gamma(\chi_{c0}(1P) \rightarrow \overline{p}n\pi^+\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ = $(2.16 \pm 0.07 \pm 0.16) \times 10^{-4}$ which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))$ = $(9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(\Lambda\overline{\Lambda})/\Gamma_{\text{total}}$	$\Gamma_{69}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID
<b><math>3.21 \pm 0.25</math> OUR FIT</b>	

$(\Lambda\overline{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{70}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>115 \pm 12 \pm 3</math></b>		426	<sup>1</sup> ABLIKIM	12i	BES3 $\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<400	90		<sup>2</sup> ABLIKIM	06D	BES2 $\psi(2S) \rightarrow \chi_{c0}\gamma$
<sup>1</sup> ABLIKIM 12i reports $(119.0 \pm 6.4 \pm 11.4) \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow \Lambda\overline{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using $B(\psi(2S) \rightarrow \chi_{c0}\gamma) = (9.2 \pm 0.5)\%$					

$(\Lambda\overline{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}$					$\Gamma_{71}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;50</b>	90	<sup>1</sup> ABLIKIM	12i	BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$	
<sup>1</sup> ABLIKIM 12i reports $< 54 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow \Lambda\overline{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .						

$\Gamma(\Sigma(1385)^+\bar{\Lambda}\pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{72}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;50</b>	90	<sup>1</sup> ABLIKIM	12i	BES3	$\psi(2S) \rightarrow \gamma \Sigma(1385)^+\bar{\Lambda}\pi^-$	
<sup>1</sup> ABLIKIM 12i reports $< 55 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{C0}(1P) \rightarrow \Sigma(1385)^+\bar{\Lambda}\pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{C0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{C0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{C0}(1P)) = 9.99 \times 10^{-2}$ .						

$\Gamma(\Sigma(1385)^-\bar{\Lambda}\pi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{73}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;50</b>	90	<sup>1</sup> ABLIKIM	12i	BES3 $\psi(2S) \rightarrow \gamma \Sigma(1385)^-\bar{\Lambda}\pi^+$	
<sup>1</sup> ABLIKIM 12i reports $< 50 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{C0}(1P) \rightarrow \Sigma(1385)^-\bar{\Lambda}\pi^+ + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{C0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{C0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{C0}(1P)) = 9.99 \times 10^{-2}$ .					

$\Gamma(K^+\overline{p}\Lambda + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_{74}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.22 \pm 0.12</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.		
$1.28 \pm 0.09 \pm 0.03$	9k	<sup>1,2</sup> ABLIKIM	13D BES3	$\psi(2S) \rightarrow \gamma \Lambda \overline{p} K^+$
$0.99 \pm 0.19 \pm 0.03$		<sup>3</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$
<sup>1</sup> ABLIKIM 13D reports $(1.32 \pm 0.03 \pm 0.10) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow K^+ \overline{p} \Lambda + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> Using $B(\Lambda \rightarrow p \pi^-) = 63.9\%$ .				
<sup>3</sup> ATHAR 07 reports $(1.07 \pm 0.17 \pm 0.12) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow K^+ \overline{p} \Lambda + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(K^+\overline{p}\Lambda(1520) + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_{75}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.9 \pm 0.7 \pm 0.1</math></b>	$62 \pm 12$	<sup>1</sup> ABLIKIM	11F	BES3 $\psi(2S) \rightarrow \gamma p \overline{p} K^+ K^-$
<sup>1</sup> ABLIKIM 11F reports $(3.00 \pm 0.58 \pm 0.50) \times 10^{-4}$ from a measurement of $\Gamma(\chi_{c0}(1P) \rightarrow K^+\overline{p}\Lambda(1520) + \text{c.c.})/\Gamma_{\text{total}} \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(\Lambda(1520)\overline{\Lambda}(1520))/\Gamma_{\text{total}}$				$\Gamma_{76}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.1 \pm 1.2 \pm 0.1</math></b>	$28 \pm 10$	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p \overline{p} K^+ K^-$
<sup>1</sup> ABLIKIM 11F reports $(3.18 \pm 1.11 \pm 0.53) \times 10^{-4}$ from a measurement of $[\Gamma(\chi_{c0}(1P) \rightarrow \Lambda(1520)\overline{\Lambda}(1520))/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(\Sigma^0\overline{\Sigma}^0)/\Gamma_{\text{total}}$				$\Gamma_{77}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.4 \pm 0.4</math> OUR AVERAGE</b>				
$4.6 \pm 0.5 \pm 0.1$	243	<sup>1</sup> ABLIKIM	13H	BES3 $\psi(2S) \rightarrow \gamma \Sigma^0 \overline{\Sigma}^0$
$4.1 \pm 0.6 \pm 0.1$	$78 \pm 10$	<sup>2</sup> NAIK	08	CLEO $\psi(2S) \rightarrow \gamma \Sigma^0 \overline{\Sigma}^0$

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# Meson Particle Listings

## $\chi_{c0}(1P)$

<sup>1</sup> ABLIKIM 13h reports  $(4.78 \pm 0.34 \pm 0.39) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Sigma^0 \bar{\Sigma}^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> NAIK 08 reports  $(4.41 \pm 0.56 \pm 0.47) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Sigma^0 \bar{\Sigma}^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Sigma^+ \bar{\Sigma}^-)/\Gamma_{\text{total}}$	$\Gamma_{78}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>3.9 \pm 0.7</math> OUR AVERAGE</b>	Error includes scale factor of 1.7.
$4.4 \pm 0.5 \pm 0.1$	148 <sup>1</sup> ABLIKIM 13h BES3 $\psi(2S) \rightarrow \gamma \Sigma^+ \bar{\Sigma}^-$
$3.0 \pm 0.6 \pm 0.1$	39 $\pm$ 7 <sup>2</sup> NAIK 08 CLEO $\psi(2S) \rightarrow \gamma \Sigma^+ \bar{\Sigma}^-$

<sup>1</sup> ABLIKIM 13h reports  $(4.54 \pm 0.42 \pm 0.30) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Sigma^+ \bar{\Sigma}^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> NAIK 08 reports  $(3.25 \pm 0.57 \pm 0.43) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Sigma^+ \bar{\Sigma}^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Sigma(1385)^+ \bar{\Sigma}(1385)^-)/\Gamma_{\text{total}}$	$\Gamma_{79}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS
<b><math>15.9 \pm 5.7 \pm 0.4</math></b>	27 <sup>1</sup> ABLIKIM 12i BES3 $\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda} \pi^+ \pi^-$

<sup>1</sup> ABLIKIM 12i reports  $(16.4 \pm 5.7 \pm 1.6) \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Sigma(1385)^+ \bar{\Sigma}(1385)^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Sigma(1385)^- \bar{\Sigma}(1385)^+)/\Gamma_{\text{total}}$	$\Gamma_{80}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS
<b><math>23 \pm 6 \pm 1</math></b>	33 <sup>1</sup> ABLIKIM 12i BES3 $\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda} \pi^+ \pi^-$

<sup>1</sup> ABLIKIM 12i reports  $(23.5 \pm 6.2 \pm 2.3) \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Sigma(1385)^- \bar{\Sigma}(1385)^+)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.68 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^- \Lambda \bar{\Xi}^+ + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{81}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>1.90 \pm 0.34 \pm 0.05</math></b>	57 <sup>1</sup> ABLIKIM 15i BES3 $\psi(2S) \rightarrow \gamma K^- \Lambda \bar{\Xi}^+ + \text{c.c.}$

<sup>1</sup> ABLIKIM 15i reports  $[(\Gamma(\chi_{c0}(1P) \rightarrow K^- \Lambda \bar{\Xi}^+ + \text{c.c.})/\Gamma_{\text{total}}) \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]] = (1.90 \pm 0.30 \pm 0.16) \times 10^{-5}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Xi^0 \bar{\Xi}^0)/\Gamma_{\text{total}}$	$\Gamma_{82}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>3.1 \pm 0.8 \pm 0.1</math></b>	23.3 $\pm$ 4.9 <sup>1</sup> NAIK 08 CLEO $\psi(2S) \rightarrow \gamma \Xi^0 \bar{\Xi}^0$

<sup>1</sup> NAIK 08 reports  $(3.34 \pm 0.70 \pm 0.48) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Xi^0 \bar{\Xi}^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Xi^- \bar{\Xi}^+)/\Gamma_{\text{total}}$	$\Gamma_{83}/\Gamma$
VALUE (units $10^{-4}$ )	CL%
<b><math>4.7 \pm 0.7 \pm 0.1</math></b>	95 $\pm$ 11 <sup>1</sup> NAIK 08 CLEO $\psi(2S) \rightarrow \gamma \Xi^- \bar{\Xi}^+$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 $< 10.3$  90 <sup>2</sup> ABLIKIM 06D BES2  $\psi(2S) \rightarrow \chi_{c0} \gamma$

<sup>1</sup> NAIK 08 reports  $(5.14 \pm 0.60 \pm 0.47) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \Xi^- \bar{\Xi}^+)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.22 \pm 0.11 \pm 0.46) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Using  $B(\psi(2S) \rightarrow \chi_{c0} \gamma) = (9.2 \pm 0.5)\%$

$\Gamma(\eta_c \pi^+ \pi^-)/\Gamma_{\text{total}}$	$\Gamma_{84}/\Gamma$
VALUE	CL%
<b><math>&lt; 7 \times 10^{-4}</math></b>	90 <sup>1,2</sup> ABLIKIM 13b BES3 $e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c0}$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 $< 41 \times 10^{-4}$  90 <sup>1,3</sup> ABLIKIM 13b BES3  $e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c0}$

<sup>1</sup> Using  $1.06 \times 10^8 \psi(2S)$  mesons and  $B(\psi(2S) \rightarrow \chi_{c0} \gamma) = (9.68 \pm 0.31)\%$ .

<sup>2</sup> From the  $\eta_c \rightarrow K_S^0 K^\pm \pi^\mp$  decays.

<sup>3</sup> From the  $\eta_c \rightarrow K^\pm K^- \pi^0$  decays.

$\Gamma(p\bar{p})/\Gamma_{\text{total}} \times \Gamma(\pi\pi)/\Gamma_{\text{total}}$	$\Gamma_{56}/\Gamma \times \Gamma_{32}/\Gamma$
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>18.8 \pm 1.2</math> OUR FIT</b>	
<b><math>15.3 \pm 2.4 \pm 0.8</math></b>	<sup>1</sup> ANDREOTTI 03 E835 $\bar{p} p \rightarrow \chi_{c0} \rightarrow \pi^0 \pi^0$

<sup>1</sup> We have multiplied  $B(p\bar{p}) \cdot B(\pi^0 \pi^0)$  measurement by 3 to obtain  $B(p\bar{p}) \cdot B(\pi\pi)$ .

$\Gamma(p\bar{p})/\Gamma_{\text{total}} \times \Gamma(\pi^0 \eta)/\Gamma_{\text{total}}$	$\Gamma_{56}/\Gamma \times \Gamma_{33}/\Gamma$
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>&lt; 0.4</math></b>	ANDREOTTI 05c E835 $\bar{p} p \rightarrow \pi^0 \eta$

$\Gamma(p\bar{p})/\Gamma_{\text{total}} \times \Gamma(\pi^0 \eta')/\Gamma_{\text{total}}$	$\Gamma_{56}/\Gamma \times \Gamma_{34}/\Gamma$
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>&lt; 2.5</math></b>	ANDREOTTI 05c E835 $\bar{p} p \rightarrow \pi^0 \eta$

$\Gamma(p\bar{p})/\Gamma_{\text{total}} \times \Gamma(\eta\eta)/\Gamma_{\text{total}}$	$\Gamma_{56}/\Gamma \times \Gamma_{36}/\Gamma$
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>6.6 \pm 0.5</math> OUR FIT</b>	
<b><math>4.0 \pm 1.2 \pm 0.5</math></b>	ANDREOTTI 05c E835 $\bar{p} p \rightarrow \eta\eta$

$\Gamma(p\bar{p})/\Gamma_{\text{total}} \times \Gamma(\eta\eta')/\Gamma_{\text{total}}$	$\Gamma_{56}/\Gamma \times \Gamma_{37}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$2.1 \pm 2.3$	ANDREOTTI 05c E835 $\bar{p} p \rightarrow \pi^0 \eta$

## RADIATIVE DECAYS

$\Gamma(\gamma J/\psi(1S))/\Gamma_{\text{total}}$	$\Gamma_{85}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID
<b><math>127 \pm 6</math> OUR FIT</b>	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$200 \pm 20 \pm 20$  <sup>1</sup> ADAM 05A CLEO  $e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c0}$

<sup>1</sup> Uses  $B(\psi(2S) \rightarrow \gamma \chi_{c0} \rightarrow \gamma \gamma J/\psi)$  from ADAM 05A and  $B(\psi(2S) \rightarrow \gamma \chi_{c0})$  from ATHAR 04.

$\Gamma(\gamma \rho^0)/\Gamma_{\text{total}}$	$\Gamma_{86}/\Gamma$
VALUE (units $10^{-6}$ )	CL%
<b><math>&lt; 9</math></b>	90 1.2 $\pm$ 4.5 <sup>1</sup> BENNETT 08A CLEO $\psi(2S) \rightarrow \gamma \gamma \rho^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 10$  90 6  $\pm$  12 <sup>2</sup> ABLIKIM 11E BES3  $\psi(2S) \rightarrow \gamma \gamma \rho^0$

<sup>1</sup> BENNETT 08A reports  $< 9.6 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \rho^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .

<sup>2</sup> ABLIKIM 11E reports  $< 10.5 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \rho^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .

$\Gamma(\gamma \omega)/\Gamma_{\text{total}}$	$\Gamma_{87}/\Gamma$
VALUE (units $10^{-6}$ )	CL%
<b><math>&lt; 8</math></b>	90 0.0 $\pm$ 2.8 <sup>1</sup> BENNETT 08A CLEO $\psi(2S) \rightarrow \gamma \gamma \omega$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 12$  90 5  $\pm$  11 <sup>2</sup> ABLIKIM 11E BES3  $\psi(2S) \rightarrow \gamma \gamma \omega$

<sup>1</sup> BENNETT 08A reports  $< 8.8 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \omega)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .

<sup>2</sup> ABLIKIM 11E reports  $< 12.9 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \omega)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .

$\Gamma(\gamma \phi)/\Gamma_{\text{total}}$	$\Gamma_{88}/\Gamma$
VALUE (units $10^{-6}$ )	CL%
<b><math>&lt; 6</math></b>	90 0.1 $\pm$ 1.6 <sup>1</sup> BENNETT 08A CLEO $\psi(2S) \rightarrow \gamma \gamma \phi$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 16$  90 15  $\pm$  7 <sup>2</sup> ABLIKIM 11E BES3  $\psi(2S) \rightarrow \gamma \gamma \phi$

<sup>1</sup> BENNETT 08A reports  $< 6.4 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \phi)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .

<sup>2</sup> ABLIKIM 11E reports  $< 16.2 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \phi)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = (9.62 \pm 0.31) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma \chi_{c0}(1P)) = 9.99 \times 10^{-2}$ .

$\Gamma(\gamma \gamma)/\Gamma_{\text{total}}$	$\Gamma_{89}/\Gamma$
VALUE (units $10^{-4}$ )	CL%
<b><math>2.23 \pm 0.13</math> OUR FIT</b>	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 7$  90 <sup>1</sup> WICHT 08 BELL  $B^\pm \rightarrow K^\pm \gamma \gamma$

<sup>1</sup> WICHT 08 reports  $[\Gamma(\chi_{c0}(1P) \rightarrow \gamma \gamma)/\Gamma_{\text{total}}] \times [B(B^\pm \rightarrow \chi_{c0}(1P) K^\pm)] < 0.11 \times 10^{-6}$  which we divide by our best value  $B(B^\pm \rightarrow \chi_{c0}(1P) K^\pm) = 1.50 \times 10^{-4}$ .

# Meson Particle Listings

## $\chi_{c0}(1P)$

### $\Gamma(\gamma\gamma)/\Gamma(\gamma J/\psi(1S))$ $\Gamma_{89}/\Gamma_{85}$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.76±0.13 OUR FIT</b>			
<b>2.0 ±0.4 OUR AVERAGE</b>			
2.2 ±0.4 <sup>+0.1</sup> <sub>-0.2</sub>	1 ANDREOTTI 04	E835	$p\bar{p} \rightarrow \chi_{c0} \rightarrow \gamma\gamma$
1.45±0.74	2 AMBROGIANI 00B	E835	$\bar{p}p \rightarrow \chi_{c2} \rightarrow \gamma\gamma, \gamma J/\psi$
1 The values of $B(p\bar{p})B(\gamma\gamma)$ and $B(\gamma\gamma)B(\gamma J/\psi)$ measured by ANDREOTTI 04 are not independent. The latter is used in the fit because of smaller systematics.			
2 Calculated by us using $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0593 \pm 0.0010$ .			

### $\Gamma(p\bar{p})/\Gamma_{total} \times \Gamma(\gamma J/\psi(1S))/\Gamma_{total}$ $\Gamma_{56}/\Gamma \times \Gamma_{85}/\Gamma$

VALUE (units $10^{-7}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>28.5±1.6 OUR FIT</b>				
<b>28.2±2.1 OUR AVERAGE</b>				
28.0±1.9±1.3	392	1,2,3 BAGNASCO 02	E835	$\bar{p}p \rightarrow \chi_{c0} \rightarrow J/\psi\gamma$
29.3 <sup>+5.7</sup> <sub>-4.7</sub> ±1.5	89	1,2 AMBROGIANI 99B		$\bar{p}p \rightarrow \chi_{c0} \rightarrow J/\psi\gamma$
1 Values in $(\Gamma(p\bar{p}) \times \Gamma(\gamma J/\psi(1S))/\Gamma_{total})$ and $(\Gamma(p\bar{p})/\Gamma_{total} \times \Gamma(\gamma J/\psi(1S))/\Gamma_{total})$ are not independent. The latter is used in the fit since it is less correlated to the total width.				
2 Calculated by us using $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0593 \pm 0.0010$ .				
3 Recalculated by ANDREOTTI 05A.				

### $\Gamma(p\bar{p})/\Gamma_{total} \times \Gamma(\gamma\gamma)/\Gamma_{total}$ $\Gamma_{56}/\Gamma \times \Gamma_{89}/\Gamma$

VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b>5.0 ±0.4 OUR FIT</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6.52±1.18 <sup>+0.48</sup> <sub>-0.72</sub>	1 ANDREOTTI 04	E835	$p\bar{p} \rightarrow \chi_{c0} \rightarrow \gamma\gamma$
1 The values of $B(p\bar{p})B(\gamma\gamma)$ and $B(\gamma\gamma)B(\gamma J/\psi)$ measured by ANDREOTTI 04 are not independent. The latter is used in the fit because of smaller systematics.			

## $\chi_{c0}(1P)$ CROSS-PARTICLE BRANCHING RATIOS

### $\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p})/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma_{total}$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>22.5±0.9 OUR FIT</b>				
<b>23.7±1.0 OUR AVERAGE</b>				
23.7±0.8±0.9	1222	ABLIKIM 13v	BES3	$\psi(2S) \rightarrow \gamma p\bar{p}$
23.7±1.4±1.4	383±22	1 NAIK 08	CLEO	$\psi(2S) \rightarrow \gamma p\bar{p}$
23.6 <sup>+3.7</sup> <sub>-3.4</sub> ±3.4	89.5 <sup>+14</sup> <sub>-13</sub>	BAI 04F	BES	$\psi(2S) \rightarrow \gamma\chi_{c0}(1P) \rightarrow \gamma\bar{p}p$

1 Calculated by us. NAIK 08 reports  $B(\chi_{c0} \rightarrow p\bar{p}) = (25.7 \pm 1.5 \pm 1.5 \pm 1.3) \times 10^{-5}$  using  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.22 \pm 0.11 \pm 0.46)\%$ .

### $\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p})/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-)$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.53±0.27 OUR FIT</b>			
<b>4.6 ±1.9</b>	1 BAI 98i	BES	$\psi(2S) \rightarrow \gamma\chi_{c0} \rightarrow \gamma\bar{p}p$

1 Calculated by us. The value for  $B(\chi_{c0} \rightarrow p\bar{p})$  reported in BAI 98i is derived using  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.3 \pm 0.8)\%$  and  $B(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = (32.4 \pm 2.6)\%$  [BAI 98D].

### $\Gamma(\chi_{c0}(1P) \rightarrow \Lambda\bar{\Lambda})/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma_{total}$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>32.0±2.3 OUR FIT</b>				
<b>31.7±2.3 OUR AVERAGE</b>				
32.0±1.9±2.2	369	1 ABLIKIM 13H	BES3	$\psi(2S) \rightarrow \gamma\Lambda\bar{\Lambda}$
31.2±3.3±2.0	131 ± 12	2 NAIK 08	CLEO	$\psi(2S) \rightarrow \gamma\Lambda\bar{\Lambda}$

1 Calculated by us. ABLIKIM 13H reports  $B(\chi_{c0} \rightarrow \Lambda\bar{\Lambda}) = (33.3 \pm 2.0 \pm 2.6) \times 10^{-5}$  from a measurement of  $B(\chi_{c0} \rightarrow \Lambda\bar{\Lambda}) \times B(\psi(2S) \rightarrow \gamma\chi_{c0})$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.62 \pm 0.31)\%$ .

2 Calculated by us. NAIK 08 reports  $B(\chi_{c0} \rightarrow \Lambda\bar{\Lambda}) = (33.8 \pm 3.6 \pm 2.2 \pm 1.7) \times 10^{-5}$  using  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.22 \pm 0.11 \pm 0.46)\%$ .

### $\Gamma(\chi_{c0}(1P) \rightarrow \Lambda\bar{\Lambda})/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-)$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.3±0.7 OUR FIT</b>				
<b>13.0<sup>+3.6</sup><sub>-3.5</sub>±2.5</b>	15.2 <sup>+4.2</sup> <sub>-4.0</sub>	1 BAI 03E	BES	$\psi(2S) \rightarrow \gamma\Lambda\bar{\Lambda}$

1 BAI 03E reports  $[B(\chi_{c0} \rightarrow \Lambda\bar{\Lambda})B(\psi(2S) \rightarrow \gamma\chi_{c0}) / B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-)] \times [B^2(\Lambda \rightarrow \pi^-p) / B(J/\psi \rightarrow p\bar{p})] = (2.45 \pm 0.68 \pm 0.65 \pm 0.46)\%$ . We calculate from this measurement the presented value using  $B(\Lambda \rightarrow \pi^-p) = (63.9 \pm 0.5)\%$  and  $B(J/\psi \rightarrow p\bar{p}) = (2.17 \pm 0.07) \times 10^{-3}$ .

### $\Gamma(\chi_{c0}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma_{total}$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.127±0.006 OUR FIT</b>				
<b>0.131±0.035 OUR AVERAGE</b>				
0.151±0.003±0.010	4.3k	ABLIKIM 12o	BES3	$\psi(2S) \rightarrow \gamma\chi_{c0}$
0.069±0.018		1 OREGLIA 82	CBAL	$\psi(2S) \rightarrow \gamma\chi_{c0}$
0.4 ±0.3		2 BRANDELIK 79B	DASP	$\psi(2S) \rightarrow \gamma\chi_{c0}$
0.16 ±0.11		2 BARTEL 78B	CNTR	$\psi(2S) \rightarrow \gamma\chi_{c0}$
3.3 ±1.7		3 BIDDICK 77	CNTR	$e^+e^- \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.125±0.007±0.013	560	4 MENDEZ 08	CLEO	$\psi(2S) \rightarrow \gamma\chi_{c0}$
0.18 ±0.01 ±0.02	172	5 ADAM 05A	CLEO	Repl. by MENDEZ 08

1 Recalculated by us using  $B(J/\psi(1S) \rightarrow \ell^+\ell^-) = 0.1181 \pm 0.0020$ .

2 Recalculated by us using  $B(J/\psi(1S) \rightarrow \mu^+\mu^-) = 0.0588 \pm 0.0010$ .

3 Assumes isotropic gamma distribution.

4 Not independent from other measurements of MENDEZ 08.

5 Not independent from other values reported by ADAM 05A.

### $\Gamma(\chi_{c0}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S)\text{anything})$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.358±0.020±0.037</b>				
0.31 ±0.02 ±0.03	172	ADAM 05A	CLEO	Repl. by MENDEZ 08

1 Not independent from other measurements of MENDEZ 08.

### $\Gamma(\chi_{c0}(1P) \rightarrow p\bar{p})/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-)$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.207±0.011 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.201±0.011±0.021	560	1 MENDEZ 08	CLEO	$\psi(2S) \rightarrow \gamma\chi_{c0}$
0.31 ±0.02 ±0.03	172	ADAM 05A	CLEO	Repl. by MENDEZ 08

1 Not independent from other measurements of MENDEZ 08.

### $\Gamma(\chi_{c0}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-)$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.367±0.019 OUR FIT</b>				
<b>0.358±0.020±0.037</b>				
0.55 ±0.04 ±0.06	172	1 ADAM 05A	CLEO	Repl. by MENDEZ 08

1 Not independent from other values reported by ADAM 05A.

### $\Gamma(\chi_{c0}(1P) \rightarrow \gamma\gamma)/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma_{total}$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.23±0.14 OUR FIT</b>				
<b>2.18±0.18 OUR AVERAGE</b>				
2.17±0.17±0.12	0.8k	ABLIKIM 12A	BES3	$\psi(2S) \rightarrow \gamma\chi_{c0} \rightarrow 3\gamma$
2.17±0.32±0.10	0.2k	ECKLUND 08A	CLEO	$\psi(2S) \rightarrow \gamma\chi_{c0} \rightarrow 3\gamma$
3.7 ±1.8 ±1.0		LEE 85	CBAL	$\psi(2S) \rightarrow \gamma\chi_{c0}$

### $\Gamma(\chi_{c0}(1P) \rightarrow \pi\pi)/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma_{total}$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.32±0.29 OUR FIT</b>				
<b>8.80±0.34 OUR AVERAGE</b>				
9.11±0.08±0.65	17k	1 ABLIKIM 10A	BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c0}$
8.81±0.11±0.43	8.9k	2 ASNER 09	CLEO	$\psi(2S) \rightarrow \gamma\pi^+\pi^-$
8.13±0.19±0.89	2.8k	3 ASNER 09	CLEO	$\psi(2S) \rightarrow \gamma\pi^0\pi^0$

1 Calculated by us. ABLIKIM 10A reports  $B(\chi_{c0} \rightarrow \pi^0\pi^0) = (3.23 \pm 0.03 \pm 0.23 \pm 0.14) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.4 \pm 0.4)\%$ . We have multiplied the  $\pi^0\pi^0$  measurement by 3 to obtain  $\pi\pi$ .

2 Calculated by us. ASNER 09 reports  $B(\chi_{c0} \rightarrow \pi^+\pi^-) = (6.37 \pm 0.08 \pm 0.31 \pm 0.32) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.22 \pm 0.11 \pm 0.46)\%$ . We have multiplied the  $\pi^+\pi^-$  measurement by 3/2 to obtain  $\pi\pi$ .

3 Calculated by us. ASNER 09 reports  $B(\chi_{c0} \rightarrow \pi^0\pi^0) = (2.94 \pm 0.07 \pm 0.32 \pm 0.15) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma\chi_{c0}) = (9.22 \pm 0.11 \pm 0.46)\%$ . We have multiplied the  $\pi^0\pi^0$  measurement by 3 to obtain  $\pi\pi$ .

### $\Gamma(\chi_{c0}(1P) \rightarrow \pi\pi)/\Gamma_{total} \times \Gamma(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-)$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>24.1±0.8 OUR FIT</b>				
<b>20.7±1.7 OUR AVERAGE</b>				
23.9±2.7±4.1	97 ± 11	1 BAI 03c	BES	$\psi(2S) \rightarrow \gamma\chi_{c0} \rightarrow \gamma\pi^0\pi^0$
20.2±1.1±1.5	720 ± 32	2 BAI 98i	BES	$\psi(2S) \rightarrow \gamma\chi_{c0} \rightarrow \gamma\pi^+\pi^-$

1 We have multiplied  $\pi^0\pi^0$  measurement by 3 to obtain  $\pi\pi$ .

2 Calculated by us. The value for  $B(\chi_{c0} \rightarrow \pi^+\pi^-)$  reported in BAI 98i is derived using  $B(\psi' \rightarrow \gamma\chi_{c0}) = (9.3 \pm 0.8)\%$  and  $B(\psi' \rightarrow J/\psi\pi^+\pi^-) = (32.4 \pm 2.6)\%$  [BAI 98D]. We have multiplied  $\pi^+\pi^-$  measurement by 3/2 to obtain  $\pi\pi$ .



# Meson Particle Listings

## $\chi_{c0}(1P)$ , $\chi_{c1}(1P)$

ANDREOTTI	05A	NP	B717 34	M. Andreotti <i>et al.</i>	(FNAL E835 Collab.)
ANDREOTTI	05C	PR	D72 112002	M. Andreotti <i>et al.</i>	(FNAL E835 Collab.)
NAKAZAWA	05	PL	B615 39	H. Nakazawa <i>et al.</i>	(BELLE Collab.)
ABE	04G	PR	D70 071102	K. Abe <i>et al.</i>	(BELLE Collab.)
ABLIKIM	04G	PR	D70 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	04H	PR	D70 092003	M. Ablikim <i>et al.</i>	(BES Collab.)
ANDREOTTI	04	PL	B584 16	M. Andreotti <i>et al.</i>	(E835 Collab.)
ATHAR	04	PR	D70 112002	S.B. Athar <i>et al.</i>	(CLEO Collab.)
BAI	04F	PR	D69 092001	J.Z. Bai <i>et al.</i>	(BES Collab.)
ANDREOTTI	03	PRL	91 091801	M. Andreotti <i>et al.</i>	(FNAL E835 Collab.)
AULCHENKO	03	PL	B573 63	V.M. Aulchenko <i>et al.</i>	(KEDR Collab.)
BAI	03C	PR	D67 032004	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	03E	PR	D67 112001	K. Abe <i>et al.</i>	(BES Collab.)
ABE,K	02	PRL	89 142001	K. Abe <i>et al.</i>	(BELLE Collab.)
BAGNASCO	02	PL	B533 237	S. Bagnasco <i>et al.</i>	(FNAL E835 Collab.)
EISENSTEIN	01	PRL	87 061801	B.I. Eisenstein <i>et al.</i>	(CLEO Collab.)
AMBROGIANI	00B	PR	D62 052002	M. Ambrogiani <i>et al.</i>	(FNAL E835 Collab.)
AMBROGIANI	99B	PRL	83 2902	M. Ambrogiani <i>et al.</i>	(FNAL E835 Collab.)
BAI	99B	PR	D60 072001	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	98D	PR	D58 092006	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	98I	PRL	81 3091	J.Z. Bai <i>et al.</i>	(BES Collab.)
GAISER	86	PR	D34 711	J. Gaiser <i>et al.</i>	(Crystal Ball Collab.)
LEE	85	SLAC	282	R.A. Lee	(SLAC)
OREGLIA	82	PR	D25 2259	M.J. Oreglia <i>et al.</i>	(SLAC, CIT, HARV+)
BRANDELIK	79B	NP	B160 426	R. Brandelik <i>et al.</i>	(DASP Collab.)
BARTEL	78B	PL	79B 492	W. Bartel <i>et al.</i>	(DESY, HEIDP)
TANENBAUM	78	PR	D17 1731	W.M. Tanenbaum <i>et al.</i>	(CLEO, LBL)
Also		Private Comm.		G. Trilling	(LBL, UCB)
BIDDICK	77	PRL	38 1324	C.J. Biddick <i>et al.</i>	(UCSD, UMD, PAVI+)

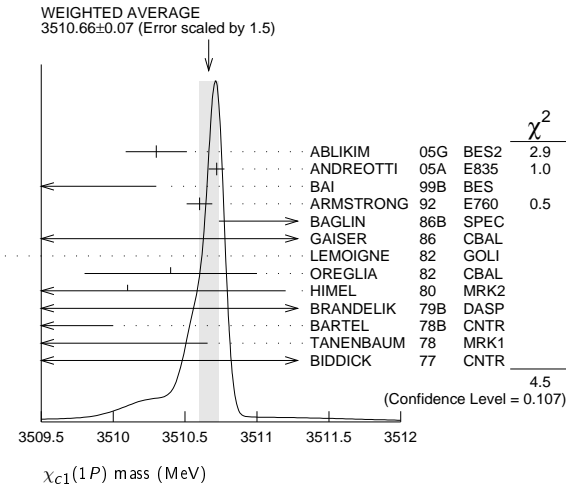
## $\chi_{c1}(1P)$

See the Review on “ $\psi(2S)$  and  $\chi_c$  branching ratios” before the  $\chi_{c0}(1P)$  Listings.

### $\chi_{c1}(1P)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3510.66 ± 0.07 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.		
3510.30 ± 0.14 ± 0.16		ABLIKIM	05G BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$
3510.719 ± 0.051 ± 0.019		ANDREOTTI	05A E835	$p\bar{p} \rightarrow e^+e^-\gamma$
3509.4 ± 0.9		BAI	99B BES	$\psi(2S) \rightarrow \gamma X$
3510.60 ± 0.087 ± 0.019	513	<sup>1</sup> ARMSTRONG	92 E760	$\bar{p}p \rightarrow e^+e^-\gamma$
3511.3 ± 0.4 ± 0.4	30	BAGLIN	86B SPEC	$\bar{p}p \rightarrow e^+e^-X$
3512.3 ± 0.3 ± 4.0		<sup>2</sup> GAISER	86 CBAL	$\psi(2S) \rightarrow \gamma X$
3507.4 ± 1.7	91	<sup>3</sup> LEMOIGNE	82 GOLI	$185 \pi^- \text{Be} \rightarrow \gamma \mu^+ \mu^- A$
3510.4 ± 0.6		OREGLIA	82 CBAL	$e^+e^- \rightarrow J/\psi 2\gamma$
3510.1 ± 1.1	254	<sup>4</sup> HIMEL	80 MRK2	$e^+e^- \rightarrow J/\psi 2\gamma$
3509 ± 11	21	BRANDELIK	79B DASP	$e^+e^- \rightarrow J/\psi 2\gamma$
3507 ± 3		<sup>4</sup> BARTEL	78B CNTR	$e^+e^- \rightarrow J/\psi 2\gamma$
3505.0 ± 4 ± 4		<sup>4,5</sup> TANENBAUM	78 MRK1	$e^+e^-$
3513 ± 7	367	<sup>4</sup> BIDDICK	77 CNTR	$\psi(2S) \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3500 ± 10	40	TANENBAUM	75 MRK1	Hadrons $\gamma$

- <sup>1</sup> Recalculated by ANDREOTTI 05A, using the value of  $\psi(2S)$  mass from AULCHENKO 03.  
<sup>2</sup> Using mass of  $\psi(2S) = 3686.0$  MeV.  
<sup>3</sup>  $J/\psi(1S)$  mass constrained to 3097 MeV.  
<sup>4</sup> Mass value shifted by us by amount appropriate for  $\psi(2S)$  mass = 3686 MeV and  $J/\psi(1S)$  mass = 3097 MeV.  
<sup>5</sup> From a simultaneous fit to radiative and hadronic decay channels.



### $\chi_{c1}(1P)$ WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.84 ± 0.04 OUR FIT</b>					
<b>0.88 ± 0.05 OUR AVERAGE</b>					
1.39 +0.40 -0.38 -0.77			ABLIKIM	05G BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$
0.876 ± 0.045 ± 0.026			ANDREOTTI	05A E835	$p\bar{p} \rightarrow e^+e^-\gamma$
0.87 ± 0.11 ± 0.08	513		<sup>1</sup> ARMSTRONG	92 E760	$\bar{p}p \rightarrow e^+e^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.3	95		BAGLIN	86B SPEC	$\bar{p}p \rightarrow e^+e^-X$
<3.8	90		GAISER	86 CBAL	$\psi(2S) \rightarrow \gamma X$
<sup>1</sup> Recalculated by ANDREOTTI 05A.					

### $\chi_{c1}(1P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Hadronic decays</b>		
$\Gamma_1$ $3(\pi^+\pi^-)$	( 5.8 ± 1.4 ) × 10 <sup>-3</sup>	S=1.2
$\Gamma_2$ $2(\pi^+\pi^-)$	( 7.6 ± 2.6 ) × 10 <sup>-3</sup>	
$\Gamma_3$ $\pi^+\pi^-\pi^0\pi^0$	( 1.22±0.16 ) %	
$\Gamma_4$ $\rho^+\pi^-\pi^0 + \text{c.c.}$	( 1.48±0.25 ) %	
$\Gamma_5$ $\rho^0\pi^+\pi^-$	( 3.9 ± 3.5 ) × 10 <sup>-3</sup>	
$\Gamma_6$ $4\pi^0$	( 5.5 ± 0.8 ) × 10 <sup>-4</sup>	
$\Gamma_7$ $\pi^+\pi^-K^+K^-$	( 4.5 ± 1.0 ) × 10 <sup>-3</sup>	
$\Gamma_8$ $K^+K^-\pi^0\pi^0$	( 1.14±0.28 ) × 10 <sup>-3</sup>	
$\Gamma_9$ $K^+K^-\pi^+\pi^-\pi^0$	( 1.15±0.13 ) %	
$\Gamma_{10}$ $K_S^0K^\pm\pi^\mp\pi^+\pi^-$	( 7.5 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{11}$ $K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	( 8.7 ± 1.4 ) × 10 <sup>-3</sup>	
$\Gamma_{12}$ $\rho^-K^+\bar{K}^0 + \text{c.c.}$	( 5.1 ± 1.2 ) × 10 <sup>-3</sup>	
$\Gamma_{13}$ $K^*(892)^0\bar{K}^0\pi^0 \rightarrow K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.}$	( 2.4 ± 0.7 ) × 10 <sup>-3</sup>	
$\Gamma_{14}$ $K^+K^-\eta\pi^0$	( 1.14±0.35 ) × 10 <sup>-3</sup>	
$\Gamma_{15}$ $\pi^+\pi^-K_S^0K_S^0$	( 7.0 ± 3.0 ) × 10 <sup>-4</sup>	
$\Gamma_{16}$ $K^+K^-\eta$	( 3.2 ± 1.0 ) × 10 <sup>-4</sup>	
$\Gamma_{17}$ $\bar{K}^0K^+\pi^- + \text{c.c.}$	( 7.1 ± 0.6 ) × 10 <sup>-3</sup>	
$\Gamma_{18}$ $K^*(892)^0\bar{K}^0 + \text{c.c.}$	( 1.0 ± 0.4 ) × 10 <sup>-3</sup>	
$\Gamma_{19}$ $K^*(892)^+K^- + \text{c.c.}$	( 1.5 ± 0.7 ) × 10 <sup>-3</sup>	
$\Gamma_{20}$ $K_J^*(1430)^0\bar{K}^0 + \text{c.c.} \rightarrow K_S^0K^+\pi^- + \text{c.c.}$	< 8 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{21}$ $K_J^*(1430)^+K^- + \text{c.c.} \rightarrow K_S^0K^+\pi^- + \text{c.c.}$	< 2.2 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{22}$ $K^+K^-\pi^0$	( 1.85±0.25 ) × 10 <sup>-3</sup>	
$\Gamma_{23}$ $\eta\pi^+\pi^-$	( 4.9 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{24}$ $a_0(980)^+\pi^- + \text{c.c.} \rightarrow \eta\pi^+\pi^-$	( 1.8 ± 0.6 ) × 10 <sup>-3</sup>	
$\Gamma_{25}$ $f_2(1270)\eta$	( 2.7 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{26}$ $\pi^+\pi^-\eta'$	( 2.3 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{27}$ $K^+K^-\eta'(958)$	( 8.8 ± 0.9 ) × 10 <sup>-4</sup>	
$\Gamma_{28}$ $K_0^*(1430)^+K^- + \text{c.c.}$	( 6.4 +2.2 -0.7 ) × 10 <sup>-4</sup>	
$\Gamma_{29}$ $f_0(980)\eta'(958)$	( 1.6 +1.4 -0.7 ) × 10 <sup>-4</sup>	
$\Gamma_{30}$ $f_0(1710)\eta'(958)$	( 7 +7 -5 ) × 10 <sup>-5</sup>	
$\Gamma_{31}$ $f_2'(1525)\eta'(958)$	( 9 ± 6 ) × 10 <sup>-5</sup>	
$\Gamma_{32}$ $\pi^0f_0(980) \rightarrow \pi^0\pi^+\pi^-$	< 6 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{33}$ $K^+\bar{K}^*(892)^0\pi^- + \text{c.c.}$	( 3.2 ± 2.1 ) × 10 <sup>-3</sup>	
$\Gamma_{34}$ $K^*(892)^0\bar{K}^*(892)^0$	( 1.5 ± 0.4 ) × 10 <sup>-3</sup>	
$\Gamma_{35}$ $K^+K^-\bar{K}_S^0K_S^0$	< 4 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{36}$ $K^+K^-\bar{K}^+K^-$	( 5.5 ± 1.1 ) × 10 <sup>-4</sup>	
$\Gamma_{37}$ $K^+K^-\phi$	( 4.2 ± 1.6 ) × 10 <sup>-4</sup>	
$\Gamma_{38}$ $\bar{K}^0K^+\pi^-\phi + \text{c.c.}$	( 3.3 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{39}$ $K^+K^-\pi^0\phi$	( 1.62±0.30 ) × 10 <sup>-3</sup>	
$\Gamma_{40}$ $\phi\pi^+\pi^-\pi^0$	( 7.5 ± 1.0 ) × 10 <sup>-4</sup>	
$\Gamma_{41}$ $\omega\omega$	( 5.8 ± 0.7 ) × 10 <sup>-4</sup>	
$\Gamma_{42}$ $\omega K^+K^-$	( 7.8 ± 0.9 ) × 10 <sup>-4</sup>	
$\Gamma_{43}$ $\omega\phi$	( 2.1 ± 0.6 ) × 10 <sup>-5</sup>	
$\Gamma_{44}$ $\phi\phi$	( 4.2 ± 0.5 ) × 10 <sup>-4</sup>	
$\Gamma_{45}$ $p\bar{p}$	( 7.72±0.35 ) × 10 <sup>-5</sup>	
$\Gamma_{46}$ $p\bar{p}\pi^0$	( 1.59±0.19 ) × 10 <sup>-4</sup>	
$\Gamma_{47}$ $p\bar{p}\eta$	( 1.48±0.25 ) × 10 <sup>-4</sup>	
$\Gamma_{48}$ $p\bar{p}\omega$	( 2.16±0.31 ) × 10 <sup>-4</sup>	
$\Gamma_{49}$ $p\bar{p}\phi$	< 1.8 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{50}$ $p\bar{p}\pi^+\pi^-$	( 5.0 ± 1.9 ) × 10 <sup>-4</sup>	
$\Gamma_{51}$ $p\bar{p}\pi^0\pi^0$		
$\Gamma_{52}$ $p\bar{p}K^+K^-$ (non-resonant)	( 1.30±0.23 ) × 10 <sup>-4</sup>	



$\Gamma_{53}$	$p\bar{p}K_S^0 K_S^0$	$< 4.5 \times 10^{-4}$	CL=90%
$\Gamma_{54}$	$p\bar{p}\pi^-$	$(3.9 \pm 0.5) \times 10^{-4}$	
$\Gamma_{55}$	$\bar{p}n\pi^+$	$(4.0 \pm 0.5) \times 10^{-4}$	
$\Gamma_{56}$	$p\bar{p}\pi^-\pi^0$	$(1.05 \pm 0.12) \times 10^{-3}$	
$\Gamma_{57}$	$\bar{p}n\pi^+\pi^0$	$(1.03 \pm 0.12) \times 10^{-3}$	
$\Gamma_{58}$	$\Lambda\bar{\Lambda}$	$(1.16 \pm 0.12) \times 10^{-4}$	
$\Gamma_{59}$	$\Lambda\bar{\Lambda}\pi^+\pi^-$	$(3.0 \pm 0.5) \times 10^{-4}$	
$\Gamma_{60}$	$\Lambda\bar{\Lambda}\pi^+\pi^-$ (non-resonant)	$(2.5 \pm 0.6) \times 10^{-4}$	
$\Gamma_{61}$	$\Sigma(1385)^+\bar{\Lambda}\pi^- + \text{c.c.}$	$< 1.3 \times 10^{-4}$	CL=90%
$\Gamma_{62}$	$\Sigma(1385)^-\bar{\Lambda}\pi^+ + \text{c.c.}$	$< 1.3 \times 10^{-4}$	CL=90%
$\Gamma_{63}$	$K^+\bar{p}\Lambda$	$(4.2 \pm 0.4) \times 10^{-4}$	S=1.1
$\Gamma_{64}$	$K^+\bar{p}\Lambda(1520) + \text{c.c.}$	$(1.7 \pm 0.5) \times 10^{-4}$	
$\Gamma_{65}$	$\Lambda(1520)\bar{\Lambda}(1520)$	$< 1.0 \times 10^{-4}$	CL=90%
$\Gamma_{66}$	$\Sigma^0\bar{\Sigma}^0$	$< 4 \times 10^{-5}$	CL=90%
$\Gamma_{67}$	$\Sigma^+\bar{\Sigma}^-$	$< 6 \times 10^{-5}$	CL=90%
$\Gamma_{68}$	$\Sigma(1385)^+\bar{\Sigma}(1385)^-$	$< 1.0 \times 10^{-4}$	CL=90%
$\Gamma_{69}$	$\Sigma(1385)^-\bar{\Sigma}(1385)^+$	$< 5 \times 10^{-5}$	CL=90%
$\Gamma_{70}$	$K^-\Lambda\Xi^+ + \text{c.c.}$	$(1.38 \pm 0.25) \times 10^{-4}$	
$\Gamma_{71}$	$\Xi^0\bar{\Xi}^0$	$< 6 \times 10^{-5}$	CL=90%
$\Gamma_{72}$	$\Xi^-\bar{\Xi}^+$	$(8.2 \pm 2.2) \times 10^{-5}$	
$\Gamma_{73}$	$\pi^+\pi^- + K^+K^-$	$< 2.1 \times 10^{-3}$	
$\Gamma_{74}$	$K_S^0 K_S^0$	$< 6 \times 10^{-5}$	CL=90%
$\Gamma_{75}$	$\eta_c\pi^+\pi^-$	$< 3.2 \times 10^{-3}$	CL=90%

## Radiative decays

$\Gamma_{76}$	$\gamma J/\psi(1S)$	$(33.9 \pm 1.2) \%$
$\Gamma_{77}$	$\gamma\rho^0$	$(2.20 \pm 0.18) \times 10^{-4}$
$\Gamma_{78}$	$\gamma\omega$	$(6.9 \pm 0.8) \times 10^{-5}$
$\Gamma_{79}$	$\gamma\phi$	$(2.5 \pm 0.5) \times 10^{-5}$
$\Gamma_{80}$	$\gamma\gamma$	

## CONSTRAINED FIT INFORMATION

A multiparticle fit to  $\chi_{c1}(1P)$ ,  $\chi_{c0}(1P)$ ,  $\chi_{c2}(1P)$ , and  $\psi(2S)$  with 4 total widths, a partial width, 25 combinations of partial widths obtained from integrated cross section, and 84 branching ratios uses 240 measurements to determine 49 parameters. The overall fit has a  $\chi^2 = 342.4$  for 191 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle\delta p_i \delta p_j\rangle/(\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ .

$x_{36}$	6				
$x_{45}$	8	3			
$x_{58}$	13	5	7		
$x_{76}$	31	13	6	26	
$\Gamma$	-19	-8	-62	-16	-51
	$x_{17}$	$x_{36}$	$x_{45}$	$x_{58}$	$x_{76}$

 $\chi_{c1}(1P)$  PARTIAL WIDTHS

$$\chi_{c1}(1P) \Gamma(i) \Gamma(\gamma J/\psi(1S))/\Gamma(\text{total})$$

$\Gamma(p\bar{p}) \times \Gamma(\gamma J/\psi(1S))/\Gamma_{\text{total}}$	$\Gamma_{45} \Gamma_{76}/\Gamma$
VALUE (eV)	DOCUMENT ID TECN COMMENT
<b>21.9±0.8 OUR FIT</b>	
<b>21.4±0.9 OUR AVERAGE</b>	
21.5±0.5±0.8	<sup>1</sup> ANDREOTTI 05A E835 $p\bar{p} \rightarrow e^+e^-\gamma$
21.4±1.5±2.2	<sup>1,2</sup> ARMSTRONG 92 E760 $\bar{p}p \rightarrow e^+e^-\gamma$
19.9±4.4	<sup>1</sup> BAGLIN 86B SPEC $\bar{p}p \rightarrow e^+e^-\chi$
19.9±4.0	

<sup>1</sup> Calculated by us using  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0593 \pm 0.0010$ .

<sup>2</sup> Recalculated by ANDREOTTI 05A.

 $\chi_{c1}(1P)$  BRANCHING RATIOS

## HADRONIC DECAYS

$\Gamma(3(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID TECN COMMENT
<b>5.8±1.4 OUR EVALUATION</b>	Error includes scale factor of 1.2. Treating systematic error as correlated.
<b>5.8±1.1 OUR AVERAGE</b>	
5.4±0.7±0.9	<sup>1</sup> BAI 99B BES $\psi(2S) \rightarrow \gamma\chi_{c1}$
16.0±5.9±0.8	<sup>1</sup> TANENBAUM 78 MRK1 $\psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> Rescaled by us using $B(\psi(2S) \rightarrow \gamma\chi_{c1}) = (8.8 \pm 0.4)\%$ and $B(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = (32.6 \pm 0.5)\%$ .	

$\Gamma(2(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID TECN COMMENT
<b>7.6±2.6 OUR EVALUATION</b>	Treating systematic error as correlated.
<b>8 ±4 OUR AVERAGE</b>	Error includes scale factor of 1.5.
4.6±2.1±2.6	<sup>1</sup> BAI 99B BES $\psi(2S) \rightarrow \gamma\chi_{c1}$
12.5±4.2±0.6	<sup>1</sup> TANENBAUM 78 MRK1 $\psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> Rescaled by us using $B(\psi(2S) \rightarrow \gamma\chi_{c1}) = (8.8 \pm 0.4)\%$ and $B(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = (32.6 \pm 0.5)\%$ .	

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE (%)	EVTS DOCUMENT ID TECN COMMENT
<b>1.22±0.15±0.04</b>	604.7 <sup>1</sup> HE 08B CLEO $e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$
<sup>1</sup> HE 08B reports $1.28 \pm 0.06 \pm 0.15 \pm 0.08 \%$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \pi^+\pi^-\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(\rho^+\pi^-\pi^0 + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
VALUE (%)	EVTS DOCUMENT ID TECN COMMENT
<b>1.48±0.24±0.05</b>	712.3 <sup>1,2</sup> HE 08B CLEO $e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$
<sup>1</sup> HE 08B reports $1.56 \pm 0.13 \pm 0.22 \pm 0.10 \%$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \rho^+\pi^-\pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	
<sup>2</sup> Calculated by us. We have added the values from HE 08B for $\rho^+\pi^-\pi^0$ and $\rho^-\pi^+\pi^0$ decays assuming uncorrelated statistical and fully correlated systematic uncertainties.	

$\Gamma(\rho^0\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b>39±35</b>	<sup>1</sup> TANENBAUM 78 MRK1 $\psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> Estimated using $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = 0.087$ . The errors do not contain the uncertainty in the $\psi(2S)$ decay.	

$\Gamma(4\pi^0)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$
VALUE (units $10^{-3}$ )	EVTS DOCUMENT ID TECN COMMENT
<b>0.55±0.08±0.02</b>	608 <sup>1</sup> ABLIKIM 11A BES3 $e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> ABLIKIM 11A reports $(0.57 \pm 0.03 \pm 0.08) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow 4\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(\pi^+\pi^-K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID TECN COMMENT
<b>4.5±1.0 OUR EVALUATION</b>	Treating systematic error as correlated.
<b>4.5±0.9 OUR AVERAGE</b>	
4.2±0.4±0.9	<sup>1</sup> BAI 99B BES $\psi(2S) \rightarrow \gamma\chi_{c1}$
7.3±3.0±0.4	<sup>1</sup> TANENBAUM 78 MRK1 $\psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> Rescaled by us using $B(\psi(2S) \rightarrow \gamma\chi_{c1}) = (8.8 \pm 0.4)\%$ and $B(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = (32.6 \pm 0.5)\%$ .	

$\Gamma(K^+K^-\pi^0\pi^0)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
VALUE (%)	EVTS DOCUMENT ID TECN COMMENT
<b>0.114±0.028±0.004</b>	45.1 <sup>1</sup> HE 08B CLEO $e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$
<sup>1</sup> HE 08B reports $0.12 \pm 0.02 \pm 0.02 \pm 0.01 \%$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^+K^-\pi^0\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(K^+K^-\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
VALUE (units $10^{-3}$ )	EVTS DOCUMENT ID TECN COMMENT
<b>11.46±0.12±1.29</b>	12k <sup>1</sup> ABLIKIM 13B BES3 $e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> Using $1.06 \times 10^8 \psi(2S)$ mesons and $B(\psi(2S) \rightarrow \chi_{c1}\gamma) = (9.2 \pm 0.4)\%$ .	

$\Gamma(K_S^0 K^\pm \pi^\mp \pi^\mp \pi^\mp)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS DOCUMENT ID TECN COMMENT
<b>7.52±0.11±0.79</b>	5.1k <sup>1</sup> ABLIKIM 13B BES3 $e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c1}$
<sup>1</sup> Using $1.06 \times 10^8 \psi(2S)$ mesons and $B(\psi(2S) \rightarrow \chi_{c1}\gamma) = (9.2 \pm 0.4)\%$ .	

$\Gamma(K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$
VALUE (%)	EVTS DOCUMENT ID TECN COMMENT
<b>0.87±0.14±0.03</b>	141.3 <sup>1</sup> HE 08B CLEO $e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$
<sup>1</sup> HE 08B reports $0.92 \pm 0.09 \pm 0.11 \pm 0.06 \%$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

# Meson Particle Listings

## $\chi_{c1}(1P)$

$\Gamma(\rho^- K^+ \bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.51 \pm 0.12 \pm 0.02</math></b>	141.3	<sup>1</sup> HE	08B CLEO	$e^+ e^- \rightarrow \gamma h^+ h^- h^0 h^0$	
<sup>1</sup> HE 08B reports $0.54 \pm 0.11 \pm 0.07 \pm 0.03$ % from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \rho^- K^+ \bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(K^*(892)^0 \bar{K}^0 \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.24 \pm 0.06 \pm 0.01</math></b>	141.3	<sup>1</sup> HE	08B CLEO	$e^+ e^- \rightarrow \gamma h^+ h^- h^0 h^0$	
<sup>1</sup> HE 08B reports $0.25 \pm 0.06 \pm 0.03 \pm 0.02$ % from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^*(892)^0 \bar{K}^0 \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(K^+ K^- \eta \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{14}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.114 \pm 0.035 \pm 0.004</math></b>	141.3	<sup>1</sup> HE	08B CLEO	$e^+ e^- \rightarrow \gamma h^+ h^- h^0 h^0$	
<sup>1</sup> HE 08B reports $0.12 \pm 0.03 \pm 0.02 \pm 0.01$ % from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- \eta \pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\pi^+ \pi^- K_S^0 \bar{K}_S^0)/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>7.0 \pm 3.0 \pm 0.2</math></b>	19.8 $\pm$ 7.7	<sup>1</sup> ABLIKIM	05o BES2	$\psi(2S) \rightarrow \chi_{c1} \gamma$	
<sup>1</sup> ABLIKIM 05o reports $[\Gamma(\chi_{c1}(1P) \rightarrow \pi^+ \pi^- K_S^0 \bar{K}_S^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ $= (0.67 \pm 0.26 \pm 0.11) \times 10^{-4}$ which we divide by our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(K^+ K^- \eta)/\Gamma_{\text{total}}$					$\Gamma_{16}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>0.32 \pm 0.10 \pm 0.01</math></b>	<sup>1</sup> ATHAR 07	CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$		
<sup>1</sup> ATHAR 07 reports $(0.34 \pm 0.10 \pm 0.04) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- \eta)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 0.0907 \pm 0.0011 \pm 0.0054$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\bar{K}^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b><math>7.1 \pm 0.6</math> OUR FIT</b>					

$\Gamma(K^*(892)^0 \bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.00 \pm 0.37 \pm 0.03</math></b>	22	<sup>1</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ABLIKIM 06R reports $(1.1 \pm 0.4 \pm 0.1) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^*(892)^0 \bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(K^*(892)^+ K^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.46 \pm 0.66 \pm 0.05</math></b>	27	<sup>1</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ABLIKIM 06R reports $(1.6 \pm 0.7 \pm 0.2) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^*(892)^+ K^- + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(K_S^*(1430)^0 \bar{K}^0 + \text{c.c.} \rightarrow K_S^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 0.8</math></b>	90	<sup>1</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ABLIKIM 06R reports $< 0.9 \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K_S^*(1430)^0 \bar{K}^0 + \text{c.c.} \rightarrow K_S^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .					

$\Gamma(K_S^*(1430)^+ K^- + \text{c.c.} \rightarrow K_S^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 2.2</math></b>	90	<sup>1</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ABLIKIM 06R reports $< 2.4 \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K_S^*(1430)^+ K^- + \text{c.c.} \rightarrow K_S^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .					

$\Gamma(K^+ K^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>1.85 \pm 0.24 \pm 0.06</math></b>	<sup>1</sup> ATHAR 07	CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$		
<sup>1</sup> ATHAR 07 reports $(1.95 \pm 0.16 \pm 0.23) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 0.0907 \pm 0.0011 \pm 0.0054$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\eta \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>4.9 \pm 0.5</math> OUR AVERAGE</b>					
$4.7 \pm 0.5 \pm 0.2$		<sup>1</sup> ATHAR 07	CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$	
$5.4 \pm 0.9 \pm 0.2$	222	<sup>2</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ATHAR 07 reports $(5.0 \pm 0.3 \pm 0.5) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 0.0907 \pm 0.0011 \pm 0.0054$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> ABLIKIM 06R reports $(5.9 \pm 0.7 \pm 0.8) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(a_0(980)^+ \pi^- + \text{c.c.} \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.8 \pm 0.6 \pm 0.1</math></b>	58	<sup>1</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ABLIKIM 06R reports $(2.0 \pm 0.5 \pm 0.5) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow a_0(980)^+ \pi^- + \text{c.c.} \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(f_2(1270) \eta)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.7 \pm 0.8 \pm 0.1</math></b>	53	<sup>1</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$	
<sup>1</sup> ABLIKIM 06R reports $(3.0 \pm 0.7 \pm 0.5) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow f_2(1270) \eta)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\pi^+ \pi^- \eta')/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>2.3 \pm 0.5 \pm 0.1</math></b>	<sup>1</sup> ATHAR 07	CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$		
<sup>1</sup> ATHAR 07 reports $(2.4 \pm 0.4 \pm 0.3) \times 10^{-3}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \pi^+ \pi^- \eta')/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 0.0907 \pm 0.0011 \pm 0.0054$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(K^+ K^- \eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>8.75 \pm 0.87</math></b>	310	<sup>1</sup> ABLIKIM	14J BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$	
<sup>1</sup> Derived using $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}) = (9.2 \pm 0.4)\%$ . Uncertainty includes both statistical and systematic contributions combined in quadrature.					

$\Gamma(K_S^*(1430)^+ K^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>6.41 \pm 0.57 \pm 2.09</math> <math>-2.71</math></b>	<sup>1</sup> ABLIKIM	14J BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$		
<sup>1</sup> Normalized to $\text{B}(\chi_{c1} \rightarrow K^+ K^- \eta'(958))$ branching fraction.					

$\Gamma(f_0(980) \eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>1.65 \pm 0.47 \pm 1.32</math> <math>-0.56</math></b>	<sup>1</sup> ABLIKIM	14J BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$		
<sup>1</sup> Normalized to $\text{B}(\chi_{c1} \rightarrow K^+ K^- \eta'(958))$ branching fraction.					

$\Gamma(f_0(1710) \eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>0.71 \pm 0.22 \pm 0.68</math> <math>-0.48</math></b>	<sup>1</sup> ABLIKIM	14J BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$		
<sup>1</sup> Normalized to $\text{B}(\chi_{c1} \rightarrow K^+ K^- \eta'(958))$ branching fraction.					

$\Gamma(f_2'(1525) \eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>0.92 \pm 0.23 \pm 0.55</math> <math>-0.51</math></b>	<sup>1</sup> ABLIKIM	14J BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$		
<sup>1</sup> Normalized to $\text{B}(\chi_{c1} \rightarrow K^+ K^- \eta'(958))$ branching fraction.					

See key on page 601

## Meson Particle Listings

 $\chi_{c1}(1P)$ 

$\Gamma(\pi^0 f_0(980) \rightarrow \pi^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{32}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6 \times 10^{-6}</math></b>	90	<sup>1</sup> ABLIKIM	11D BES3	$\psi(2S) \rightarrow \gamma \pi^0 \pi^+ \pi^-$	

<sup>1</sup> ABLIKIM 11D reports  $[\Gamma(\chi_{c1}(1P) \rightarrow \pi^0 f_0(980) \rightarrow \pi^0 \pi^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))] < 6.0 \times 10^{-7}$  which we divide by our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .

$\Gamma(K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>32 \pm 21</math></b>	<sup>1</sup> TANENBAUM	78	MRK1	$\psi(2S) \rightarrow \gamma \chi_{c1}$	

<sup>1</sup> Estimated using  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 0.087$ . The errors do not contain the uncertainty in the  $\psi(2S)$  decay.

$\Gamma(K^*(892)^0 \bar{K}^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{34}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.47 \pm 0.36 \pm 0.05</math></b>	$28.4 \pm 5.5$	<sup>1,2</sup> ABLIKIM	04H BES	$\psi(2S) \rightarrow \gamma K^+ K^- \pi^+ \pi^-$	

<sup>1</sup> ABLIKIM 04H reports  $[\Gamma(\chi_{c1}(1P) \rightarrow K^*(892)^0 \bar{K}^*(892)^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))] = (1.40 \pm 0.27 \pm 0.22) \times 10^{-4}$  which we divide by our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Assumes  $\text{B}(K^*(892)^0 \rightarrow K^- \pi^+) = 2/3$ .

$\Gamma(K^+ K^- K_S^0 K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{35}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;4</math></b>	90	$3.2 \pm 2.4$	<sup>1</sup> ABLIKIM	05o BES2	$\psi(2S) \rightarrow \chi_{c1} \gamma$

<sup>1</sup> ABLIKIM 05o reports  $[\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- K_S^0 K_S^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))] < 4.2 \times 10^{-5}$  which we divide by our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .

$\Gamma(K^+ K^- K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{36}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b><math>0.55 \pm 0.11</math> OUR FIT</b>					

$\Gamma(K^+ K^- \phi)/\Gamma_{\text{total}}$					$\Gamma_{37}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.42 \pm 0.15 \pm 0.01</math></b>	17	<sup>1</sup> ABLIKIM	06T BES2	$\psi(2S) \rightarrow \gamma 2K^+ 2K^-$	

<sup>1</sup> ABLIKIM 06T reports  $(0.46 \pm 0.16 \pm 0.06) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\bar{K}^0 K^+ \pi^- \phi + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{38}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>3.27 \pm 0.28 \pm 0.46</math></b>	ABLIKIM	15M BES3	$\psi(2S) \rightarrow \gamma \chi_{c1}$		

$\Gamma(K^+ K^- \pi^0 \phi)/\Gamma_{\text{total}}$					$\Gamma_{39}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>1.62 \pm 0.12 \pm 0.28</math></b>	ABLIKIM	15M BES3	$\psi(2S) \rightarrow \gamma \chi_{c1}$		

$\Gamma(\phi \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{40}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.75 \pm 0.06 \pm 0.08</math></b>	373	<sup>1</sup> ABLIKIM	13B BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c1}$	

<sup>1</sup> Using  $1.06 \times 10^8$   $\psi(2S)$  mesons and  $\text{B}(\psi(2S) \rightarrow \chi_{c1} \gamma) = (9.2 \pm 0.4)\%$ .

$\Gamma(\omega \gamma)/\Gamma_{\text{total}}$					$\Gamma_{41}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>5.8 \pm 0.7 \pm 0.2</math></b>	597	<sup>1</sup> ABLIKIM	11K BES3	$\psi(2S) \rightarrow \gamma$ hadrons	

<sup>1</sup> ABLIKIM 11K reports  $(6.0 \pm 0.3 \pm 0.7) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow \omega \gamma)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\omega K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{42}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.78 \pm 0.04 \pm 0.08</math></b>	628	<sup>1</sup> ABLIKIM	13B BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c1}$	

<sup>1</sup> Using  $1.06 \times 10^8$   $\psi(2S)$  mesons and  $\text{B}(\psi(2S) \rightarrow \chi_{c1} \gamma) = (9.2 \pm 0.4)\%$ .

$\Gamma(\omega \phi)/\Gamma_{\text{total}}$					$\Gamma_{43}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.21 \pm 0.06 \pm 0.01</math></b>	15	<sup>1</sup> ABLIKIM	11K BES3	$\psi(2S) \rightarrow \gamma$ hadrons	

<sup>1</sup> ABLIKIM 11K reports  $(0.22 \pm 0.06 \pm 0.02) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow \omega \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\phi \phi)/\Gamma_{\text{total}}$					$\Gamma_{44}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>4.2 \pm 0.5 \pm 0.1</math></b>	366	<sup>1</sup> ABLIKIM	11K BES3	$\psi(2S) \rightarrow \gamma$ hadrons	

<sup>1</sup> ABLIKIM 11K reports  $(4.4 \pm 0.3 \pm 0.5) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow \phi \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p\bar{p})/\Gamma_{\text{total}}$		$\Gamma_{45}/\Gamma$	
VALUE (units $10^{-4}$ )	DOCUMENT ID		
<b><math>0.772 \pm 0.035</math> OUR FIT</b>			

$\Gamma(p\bar{p}\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{46}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.159 \pm 0.019</math> OUR AVERAGE</b>				

$0.166 \pm 0.020 \pm 0.005$  <sup>1</sup> ONYISI 10 CLE3  $\psi(2S) \rightarrow \gamma p\bar{p}X$

$0.114 \pm 0.048 \pm 0.004$  <sup>2</sup> ATHAR 07 CLEO  $\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ONYISI 10 reports  $(1.75 \pm 0.16 \pm 0.13 \pm 0.11) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ATHAR 07 reports  $(1.2 \pm 0.5 \pm 0.1) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p\bar{p}\eta)/\Gamma_{\text{total}}$					$\Gamma_{47}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>0.148 \pm 0.025 \pm 0.005</math></b>		<sup>1</sup> ONYISI	10 CLE3	$\psi(2S) \rightarrow \gamma p\bar{p}X$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.15$

<sup>1</sup> ONYISI 10 reports  $(1.56 \pm 0.22 \pm 0.14 \pm 0.10) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\eta)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ATHAR 07 reports  $<0.16 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\eta)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .

$\Gamma(p\bar{p}\omega)/\Gamma_{\text{total}}$					$\Gamma_{48}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>0.216 \pm 0.031 \pm 0.007</math></b>	<sup>1</sup> ONYISI	10 CLE3	$\psi(2S) \rightarrow \gamma p\bar{p}X$		

<sup>1</sup> ONYISI 10 reports  $(2.28 \pm 0.28 \pm 0.16 \pm 0.14) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p\bar{p}\phi)/\Gamma_{\text{total}}$					$\Gamma_{49}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.8</math></b>	90	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\bar{p}K^+ K^-$	

<sup>1</sup> ABLIKIM 11F reports  $<1.82 \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .

$\Gamma(p\bar{p}\pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b><math>0.50 \pm 0.19</math> OUR EVALUATION</b>	Treating systematic error as correlated.				

**$0.50 \pm 0.19$  OUR AVERAGE**

$0.46 \pm 0.12 \pm 0.15$

$1.08 \pm 0.77 \pm 0.05$

<sup>1</sup> Rescaled by us using  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}) = (8.8 \pm 0.4)\%$  and  $\text{B}(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (32.6 \pm 0.5)\%$ .

$\Gamma(p\bar{p}\pi^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{51}/\Gamma$
VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT	
$<0.05$	90	<sup>1</sup> HE	08B CLEO	$e^+ e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $<0.05\%$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}\pi^0 \pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = 9.55 \times 10^{-2}$ .

$\Gamma(p\bar{p}K^+ K^- (\text{non-resonant}))/\Gamma_{\text{total}}$					$\Gamma_{52}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.30 \pm 0.23 \pm 0.04</math></b>	$82 \pm 9$	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\bar{p}K^+ K^-$	

<sup>1</sup> ABLIKIM 11F reports  $(1.35 \pm 0.15 \pm 0.19) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow p\bar{p}K^+ K^- (\text{non-resonant}))/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

## $\chi_{c1}(1P)$

$\Gamma(p\overline{p}K_S^0K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{53}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;4.5</b>	90	<sup>1</sup> ABLIKIM	06D BES2	$\psi(2S) \rightarrow \gamma\chi_{c1}$	
<sup>1</sup> Using $B(\psi(2S) \rightarrow \chi_{c1}\gamma)$ (9.1 ± 0.6)%.					

$\Gamma(p\overline{p}\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{54}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3.9 ± 0.5 ± 0.1</b>	1412	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma p\overline{p}\pi^-$	
<sup>1</sup> ABLIKIM 12J reports $[\Gamma(\chi_{c1}(1P) \rightarrow p\overline{p}\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ = (0.37 ± 0.02 ± 0.04) × 10 <sup>-4</sup> which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\overline{p}n\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{55}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.0 ± 0.5 ± 0.1</b>	1625	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma\overline{p}n\pi^+$	
<sup>1</sup> ABLIKIM 12J reports $[\Gamma(\chi_{c1}(1P) \rightarrow \overline{p}n\pi^+)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ = (0.38 ± 0.02 ± 0.04) × 10 <sup>-4</sup> which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(p\overline{p}\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{56}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>10.5 ± 1.2 ± 0.3</b>	1082	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma p\overline{p}\pi^-\pi^0$	
<sup>1</sup> ABLIKIM 12J reports $[\Gamma(\chi_{c1}(1P) \rightarrow p\overline{p}\pi^-\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ = (1.00 ± 0.05 ± 0.10) × 10 <sup>-4</sup> which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\overline{p}n\pi^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{57}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>10.3 ± 1.2 ± 0.3</b>	1261	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma\overline{p}n\pi^+\pi^0$	
<sup>1</sup> ABLIKIM 12J reports $[\Gamma(\chi_{c1}(1P) \rightarrow \overline{p}n\pi^+\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ = (0.98 ± 0.05 ± 0.10) × 10 <sup>-4</sup> which we divide by our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\Lambda\overline{\Lambda})/\Gamma_{\text{total}}$					$\Gamma_{58}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID				
<b>1.16 ± 0.12 OUR FIT</b>					
$\Gamma(\Lambda\overline{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{59}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>30 ± 5 ± 1</b>	105		<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<150	90		<sup>2</sup> ABLIKIM	06D BES2	$\psi(2S) \rightarrow \gamma\chi_{c1}$

$\Gamma(\Lambda\overline{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}$					$\Gamma_{60}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>25 ± 6 ± 1</b>	13	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$	
<sup>1</sup> ABLIKIM 12i reports (26.2 ± 5.5 ± 3.3) × 10 <sup>-5</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Lambda\overline{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using $B(\psi(2S) \rightarrow \chi_{c1}\gamma)$ (9.1 ± 0.6)%.					

$\Gamma(\Lambda\overline{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}$					$\Gamma_{60}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>25 ± 6 ± 1</b>	13	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$	
<sup>1</sup> ABLIKIM 12i reports (26.2 ± 5.5 ± 3.3) × 10 <sup>-5</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Lambda\overline{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\Sigma(1385)^+\overline{\Lambda}\pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{61}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;13</b>	90	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Sigma(1385)^+\overline{\Lambda}\pi^-$	
<sup>1</sup> ABLIKIM 12i reports < 14 × 10 <sup>-5</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma(1385)^+\overline{\Lambda}\pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(\Sigma(1385)^-\overline{\Lambda}\pi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{62}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;13</b>	90	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Sigma(1385)^-\overline{\Lambda}\pi^+$	
<sup>1</sup> ABLIKIM 12i reports < 14 × 10 <sup>-5</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma(1385)^-\overline{\Lambda}\pi^+ + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(K^+\overline{p}\Lambda)/\Gamma_{\text{total}}$					$\Gamma_{63}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.2 ± 0.4 OUR AVERAGE</b>	Error includes scale factor of 1.1.				
4.3 ± 0.4 ± 0.1	3k	<sup>1,2</sup> ABLIKIM	13D BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{p}K^+$	
3.1 ± 0.9 ± 0.1		<sup>3</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$	

<sup>1</sup> ABLIKIM 13D reports (4.5 ± 0.2 ± 0.4) × 10<sup>-4</sup> from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow K^+\overline{p}\Lambda)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$  = (9.2 ± 0.4) × 10<sup>-2</sup>, which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$  = (9.55 ± 0.31) × 10<sup>-2</sup>. Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Using  $B(\Lambda \rightarrow p\pi^-)$  = 63.9%.

<sup>3</sup> ATHAR 07 reports (3.3 ± 0.9 ± 0.4) × 10<sup>-4</sup> from a measurement of  $[\Gamma(\chi_{c1}(1P) \rightarrow K^+\overline{p}\Lambda)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$  = (9.07 ± 0.11 ± 0.54) × 10<sup>-2</sup>, which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$  = (9.55 ± 0.31) × 10<sup>-2</sup>. Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+\overline{p}\Lambda(1520) + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{64}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.7 ± 0.4 ± 0.1</b>	48 ± 10	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\overline{p}K^+ K^-$	
<sup>1</sup> ABLIKIM 11F reports (1.81 ± 0.38 ± 0.28) × 10 <sup>-4</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow K^+\overline{p}\Lambda(1520) + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.55 ± 0.31) × 10 <sup>-2</sup> . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(\Lambda(1520)\overline{\Lambda}(1520))/\Gamma_{\text{total}}$					$\Gamma_{65}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.0</b>	90	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\overline{p}K^+ K^-$	
<sup>1</sup> ABLIKIM 11F reports < 1.00 × 10 <sup>-4</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Lambda(1520)\overline{\Lambda}(1520))/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(\Sigma^0\overline{\Sigma}^0)/\Gamma_{\text{total}}$					$\Gamma_{66}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.4</b>	90	3.8 ± 2.5	<sup>1</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma\Sigma^0\overline{\Sigma}^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.6	90		<sup>2</sup> ABLIKIM	13H BES3	$\psi(2S) \rightarrow \gamma\Sigma^0\overline{\Sigma}^0$
<sup>1</sup> NAIK 08 reports < 0.44 × 10 <sup>-4</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma^0\overline{\Sigma}^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.07 ± 0.11 ± 0.54) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					
<sup>2</sup> ABLIKIM 13H reports < 0.62 × 10 <sup>-4</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma^0\overline{\Sigma}^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(\Sigma^+\overline{\Sigma}^-)/\Gamma_{\text{total}}$					$\Gamma_{67}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.6</b>	90	4.3 ± 2.3	<sup>1</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma\Sigma^+\overline{\Sigma}^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.8	90		<sup>2</sup> ABLIKIM	13H BES3	$\psi(2S) \rightarrow \gamma\Sigma^+\overline{\Sigma}^-$
<sup>1</sup> NAIK 08 reports < 0.65 × 10 <sup>-4</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma^+\overline{\Sigma}^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.07 ± 0.11 ± 0.54) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					
<sup>2</sup> ABLIKIM 13H reports < 0.87 × 10 <sup>-4</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma^+\overline{\Sigma}^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(\Sigma(1385)^+\overline{\Sigma}(1385)^-)/\Gamma_{\text{total}}$					$\Gamma_{68}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;10</b>	90	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$	
<sup>1</sup> ABLIKIM 12i reports < 10 × 10 <sup>-5</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma(1385)^+\overline{\Sigma}(1385)^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(\Sigma(1385)^-\overline{\Sigma}(1385)^+)/\Gamma_{\text{total}}$					$\Gamma_{69}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;5</b>	90	<sup>1</sup> ABLIKIM	12i BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{\Lambda}\pi^+\pi^-$	
<sup>1</sup> ABLIKIM 12i reports < 5.7 × 10 <sup>-5</sup> from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Sigma(1385)^-\overline{\Sigma}(1385)^+)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$ assuming $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = (9.2 ± 0.4) × 10 <sup>-2</sup> , which we rescale to our best value $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))$ = 9.55 × 10 <sup>-2</sup> .					

$\Gamma(K^- \Lambda \Xi^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{70}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.38 \pm 0.24 \pm 0.05</math></b>	49	<sup>1</sup> ABLIKIM	151 BES3	$\psi(2S) \rightarrow \gamma K^- \Lambda \Xi^+ + \text{c.c.}$
<sup>1</sup> ABLIKIM 151 reports $[\Gamma(\chi_{c1}(1P) \rightarrow K^- \Lambda \Xi^+ + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ = $(1.32 \pm 0.20 \pm 0.12) \times 10^{-5}$ which we divide by our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

 $\Gamma(\Xi^0 \Xi^0)/\Gamma_{\text{total}}$   $\Gamma_{71}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.6</b>	90	1.7 $\pm$ 2.4	<sup>1</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma \Xi^0 \Xi^0$
<sup>1</sup> NAIK 08 reports $< 0.60 \times 10^{-4}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Xi^0 \Xi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $9.55 \times 10^{-2}$ .					

 $\Gamma(\Xi^- \Xi^+)/\Gamma_{\text{total}}$   $\Gamma_{72}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.82 \pm 0.22 \pm 0.03</math></b>	16.4 $\pm$ 4.3		<sup>1</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma \Xi^+ \Xi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 3.4	90		<sup>2</sup> ABLIKIM	06d BES2	$\psi(2S) \rightarrow \gamma \chi_{c1}$
<sup>1</sup> NAIK 08 reports $(0.86 \pm 0.22 \pm 0.08) \times 10^{-4}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \Xi^- \Xi^+)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.07 \pm 0.11 \pm 0.54) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using $\text{B}(\psi(2S) \rightarrow \chi_{c1} \gamma)$ (9.1 $\pm$ 0.6)%.					

 $[\Gamma(\pi^+ \pi^-) + \Gamma(K^+ K^-)]/\Gamma_{\text{total}}$   $\Gamma_{73}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;21</b>		<sup>1</sup> FELDMAN	77 MRK1	$\psi(2S) \rightarrow \gamma \chi_{c1}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<38	90	<sup>1</sup> BRANDELIK	79b DASP	$\psi(2S) \rightarrow \gamma \chi_{c1}$
<sup>1</sup> Estimated using $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = 0.087. The errors do not contain the uncertainty in the $\psi(2S)$ decay.				

 $\Gamma(K_S^0 K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{74}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.6</b>	90	<sup>1</sup> ABLIKIM	05o BES2	$\psi(2S) \rightarrow \chi_{c1} \gamma$
<sup>1</sup> ABLIKIM 05o reports $[\Gamma(\chi_{c1}(1P) \rightarrow K_S^0 K_S^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ < $0.6 \times 10^{-5}$ which we divide by our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $9.55 \times 10^{-2}$ .				

 $\Gamma(\eta_c \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{75}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.2 <math>\times 10^{-3}</math></b>	90	<sup>1,2</sup> ABLIKIM	13b BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c1}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<4.4 $\times 10^{-3}$	90	<sup>1,3</sup> ABLIKIM	13b BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c1}$
<sup>1</sup> Using $1.06 \times 10^8$ $\psi(2S)$ mesons and $\text{B}(\psi(2S) \rightarrow \chi_{c1} \gamma)$ = $(9.2 \pm 0.4)\%$ .				
<sup>2</sup> Using the $\eta_c \rightarrow K_S^0 K^\pm \pi^\mp$ decays.				
<sup>3</sup> Using the $\eta_c \rightarrow K^+ K^- \pi^0$ decays.				

## RADIATIVE DECAYS

 $\Gamma(\gamma J/\psi(1S))/\Gamma_{\text{total}}$   $\Gamma_{76}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.339 \pm 0.012</math> OUR FIT</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.379 $\pm$ 0.008 $\pm$ 0.021	<sup>1</sup> ADAM	05A CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c1}$
<sup>1</sup> Uses $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1} \rightarrow \gamma \gamma J/\psi)$ from ADAM 05A and $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1})$ from ATHAR 04.			

 $\Gamma(\gamma \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{77}/\Gamma$ 

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>220 \pm 18</math> OUR AVERAGE</b>				
220 $\pm$ 23 $\pm$ 7	432 $\pm$ 25	<sup>1</sup> ABLIKIM	11E BES3	$\psi(2S) \rightarrow \gamma \gamma \rho^0$
221 $\pm$ 24 $\pm$ 7	186 $\pm$ 15	<sup>2</sup> BENNETT	08A CLEO	$\psi(2S) \rightarrow \gamma \gamma \rho^0$
<sup>1</sup> ABLIKIM 11E reports $(228 \pm 13 \pm 22) \times 10^{-6}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \gamma \rho^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> BENNETT 08A reports $(243 \pm 19 \pm 22) \times 10^{-6}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \gamma \rho^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

 $\Gamma(\gamma \omega)/\Gamma_{\text{total}}$   $\Gamma_{78}/\Gamma$ 

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>69 \pm 8</math> OUR AVERAGE</b>				
67 $\pm$ 9 $\pm$ 2	136 $\pm$ 14	<sup>1</sup> ABLIKIM	11E BES3	$\psi(2S) \rightarrow \gamma \gamma \omega$
76 $\pm$ 17 $\pm$ 2	39 $\pm$ 7	<sup>2</sup> BENNETT	08A CLEO	$\psi(2S) \rightarrow \gamma \gamma \omega$
<sup>1</sup> ABLIKIM 11E reports $(69.7 \pm 7.2 \pm 6.6) \times 10^{-6}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \gamma \omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> BENNETT 08A reports $(83 \pm 15 \pm 12) \times 10^{-6}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \gamma \omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

 $\Gamma(\gamma \phi)/\Gamma_{\text{total}}$   $\Gamma_{79}/\Gamma$ 

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>25 \pm 5 \pm 1</math></b>	43 $\pm$ 9		<sup>1</sup> ABLIKIM	11E BES3	$\psi(2S) \rightarrow \gamma \gamma \phi$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<24	90	5.2 $\pm$ 3.1	<sup>2</sup> BENNETT	08A CLEO	$\psi(2S) \rightarrow \gamma \gamma \phi$
<sup>1</sup> ABLIKIM 11E reports $(25.8 \pm 5.2 \pm 2.3) \times 10^{-6}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \gamma \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.2 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(9.55 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> BENNETT 08A reports $< 26 \times 10^{-6}$ from a measurement of $[\Gamma(\chi_{c1}(1P) \rightarrow \gamma \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $(8.7 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = $9.55 \times 10^{-2}$ .					

 $\Gamma(\gamma \gamma)/\Gamma_{\text{total}}$   $\Gamma_{80}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 3.5	90	ECKLUND	08A CLEO	$\psi(2S) \rightarrow \gamma \chi_{c1} \rightarrow 3\gamma$
<150	90	<sup>1</sup> YAMADA	77 DASP	$e^+ e^- \rightarrow 3\gamma$
<sup>1</sup> Estimated using $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))$ = 0.087. The errors do not contain the uncertainty in the $\psi(2S)$ decay.				

 $\chi_{c1}(1P)$  CROSS-PARTICLE BRANCHING RATIOS $\Gamma(\chi_{c1}(1P) \rightarrow p \bar{p})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$   $\Gamma_{45}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$ 

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.14 \pm 0.11</math> OUR FIT</b>			
<b><math>1.1 \pm 1.0</math></b>	<sup>1</sup> BAI	98i BES	$\psi(2S) \rightarrow \gamma \chi_{c1} \rightarrow \gamma p \bar{p}$
<sup>1</sup> Calculated by us. The value for $\text{B}(\chi_{c1} \rightarrow p \bar{p})$ reported in BAI 98i is derived using $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1})$ = $(8.7 \pm 0.8)\%$ and $\text{B}(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$ = $(32.4 \pm 2.6)\%$ [BAI 98b].			

 $\Gamma(\chi_{c1}(1P) \rightarrow \Lambda \bar{\Lambda})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}}$   $\Gamma_{58}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$ 

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>11.1 \pm 1.1</math> OUR FIT</b>				
<b><math>10.9 \pm 1.1</math> OUR AVERAGE</b>				
11.2 $\pm$ 1.0 $\pm$ 0.9	136	<sup>1</sup> ABLIKIM	13H BES3	$\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda}$
10.5 $\pm$ 1.6 $\pm$ 0.6	46 $\pm$ 7	<sup>2</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda}$
<sup>1</sup> Calculated by us. ABLIKIM 13H reports $\text{B}(\chi_{c1} \rightarrow \Lambda \bar{\Lambda})$ = $(12.2 \pm 1.1 \pm 1.1) \times 10^{-5}$ from a measurement of $\text{B}(\chi_{c1} \rightarrow \Lambda \bar{\Lambda}) \times \text{B}(\psi(2S) \rightarrow \gamma \chi_{c1})$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1})$ = $(9.2 \pm 0.4)\%$ .				
<sup>2</sup> Calculated by us. NAIK 08 reports $\text{B}(\chi_{c1} \rightarrow \Lambda \bar{\Lambda})$ = $(11.6 \pm 1.8 \pm 0.7 \pm 0.7) \times 10^{-5}$ using $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c1})$ = $(9.07 \pm 0.11 \pm 0.54)\%$ .				

 $\Gamma(\chi_{c1}(1P) \rightarrow \Lambda \bar{\Lambda})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$   $\Gamma_{58}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.22 \pm 0.31</math> OUR FIT</b>				
<b><math>7.1 \pm 2.8 \pm 1.3</math></b>	9.0 $\pm$ 3.5 $\pm$ 3.1	<sup>1</sup> BAI	03E BES	$\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda}$
<sup>1</sup> BAI 03E reports $[\text{B}(\chi_{c1} \rightarrow \Lambda \bar{\Lambda}) \text{B}(\psi(2S) \rightarrow \gamma \chi_{c1}) / \text{B}(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-)] \times [\text{B}^2(\Lambda \rightarrow \pi^- p) / \text{B}(J/\psi \rightarrow p \bar{p})]$ = $(1.33 \pm 0.52 \pm 0.25)\%$ . We calculate from this measurement the presented value using $\text{B}(\Lambda \rightarrow \pi^- p)$ = $(63.9 \pm 0.5)\%$ and $\text{B}(J/\psi \rightarrow p \bar{p})$ = $(2.17 \pm 0.07) \times 10^{-3}$ .				

## Meson Particle Listings

 $\chi_{c1}(1P)$ 

$$\Gamma(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}} \times \Gamma_{76}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{133}^{\psi(2S)}$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.24 ± 0.07 OUR FIT</b>				
<b>2.93 ± 0.15 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
3.377 ± 0.009 ± 0.183	142k	ABLIKIM	120	BES3 $\psi(2S) \rightarrow \gamma \chi_{c1}$
2.81 ± 0.05 ± 0.23	13k	BAI	04i	BES2 $\psi(2S) \rightarrow J/\psi \gamma \gamma$
2.56 ± 0.12 ± 0.20		GAISER	86	CBAL $\psi(2S) \rightarrow \gamma X$
2.78 ± 0.30		<sup>1</sup> OREGLIA	82	CBAL $\psi(2S) \rightarrow \gamma \chi_{c1}$
2.2 ± 0.5		<sup>2</sup> BRANDELIK	79B	DASP $\psi(2S) \rightarrow \gamma \chi_{c1}$
2.9 ± 0.5		<sup>2</sup> BARTEL	78B	CNTR $\psi(2S) \rightarrow \gamma \chi_{c1}$
5.0 ± 1.5		<sup>3</sup> BIDDICK	77	CNTR $e^+ e^- \rightarrow \gamma X$
2.8 ± 0.9		<sup>1</sup> WHITAKER	76	MRK1 $e^+ e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.56 ± 0.03 ± 0.12	24.9k	<sup>4</sup> MENDEZ	08	CLEO $\psi(2S) \rightarrow \gamma \chi_{c1}$
3.44 ± 0.06 ± 0.13	3.7k	<sup>5</sup> ADAM	05A	CLEO Repl. by MENDEZ 08

<sup>1</sup> Recalculated by us using  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .

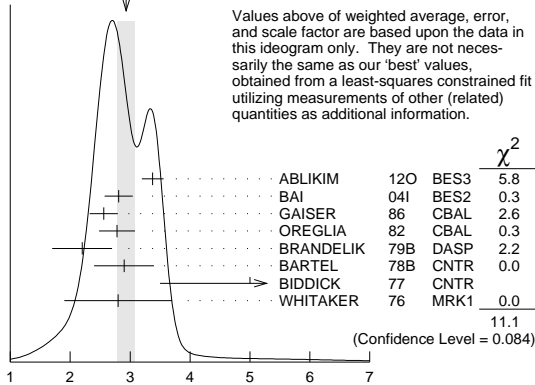
<sup>2</sup> Recalculated by us using  $B(J/\psi(1S) \rightarrow \mu^+ \mu^-) = 0.0588 \pm 0.0010$ .

<sup>3</sup> Assumes isotropic gamma distribution.

<sup>4</sup> Not independent from other measurements of MENDEZ 08.

<sup>5</sup> Not independent from other values reported by ADAM 05A.

WEIGHTED AVERAGE  
2.93±0.15 (Error scaled by 1.4)



$$\Gamma(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}} \text{ (units } 10^{-2})$$

$$\Gamma(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{76}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{133}^{\psi(2S)}$$

$$\Gamma_{76}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{133}^{\psi(2S)} = \Gamma_{76}/\Gamma \times \Gamma_{133}^{\psi(2S)}/(\Gamma_{11}^{\psi(2S)} + \Gamma_{12}^{\psi(2S)} + \Gamma_{13}^{\psi(2S)} + 0.339\Gamma_{133}^{\psi(2S)} + 0.192\Gamma_{134}^{\psi(2S)})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.31 ± 0.11 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.70 ± 0.04 ± 0.15	24.9k	<sup>1</sup> MENDEZ	08	CLEO $\psi(2S) \rightarrow \gamma \chi_{c1}$
5.77 ± 0.10 ± 0.12	3.7k	ADAM	05A	CLEO Repl. by MENDEZ 08

<sup>1</sup> Not independent from other measurements of MENDEZ 08.

$$\Gamma(\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S))/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{76}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.40 ± 0.21 OUR FIT</b>				
<b>10.15 ± 0.28 OUR AVERAGE</b>				
10.17 ± 0.07 ± 0.27	24.9k	MENDEZ	08	CLEO $\psi(2S) \rightarrow \gamma \chi_{c1}$
12.6 ± 0.3 ± 3.8	3k	<sup>1</sup> ABLIKIM	04B	BES $\psi(2S) \rightarrow J/\psi X$
8.5 ± 2.1		<sup>2</sup> HIMEL	80	MRK2 $\psi(2S) \rightarrow \gamma \chi_{c1}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
10.24 ± 0.17 ± 0.23	3.7k	<sup>3</sup> ADAM	05A	CLEO Repl. by MENDEZ 08

<sup>1</sup> From a fit to the  $J/\psi$  recoil mass spectra.

<sup>2</sup> The value for  $B(\psi(2S) \rightarrow \gamma \chi_{c1}) \times B(\chi_{c1} \rightarrow \gamma J/\psi(1S))$  quoted in HIMEL 80 is derived using  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (33 \pm 3)\%$  and  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.138 \pm 0.018$ . Calculated by us using  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .

<sup>3</sup> Not independent from other values reported by ADAM 05A.

$$\Gamma(\chi_{c1}(1P) \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}} \times \Gamma_{17}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{133}^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.8 ± 0.5 OUR FIT</b>			
<b>7.2 ± 0.6 OUR AVERAGE</b>			
7.3 ± 0.5 ± 0.5	<sup>1</sup> ATHAR	07	CLEO $\psi(2S) \rightarrow \gamma K_S^0 K^+ \pi^-$
7.0 ± 0.5 ± 0.9	<sup>2</sup> ABLIKIM	06R	BES2 $\psi(2S) \rightarrow \gamma \chi_{c1}$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c1} \rightarrow K^0 K^+ \pi^- + \text{c.c.})$ reported by ATHAR 07 was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (9.07 \pm 0.11 \pm 0.54)\%$ .			
<sup>2</sup> Calculated by us. ABLIKIM 06R reports $B(\chi_{c1} \rightarrow K_S^0 K^+ \pi^-) = (4.0 \pm 0.3 \pm 0.5) \times 10^{-3}$ . We use $B(\psi(2S) \rightarrow \gamma \chi_{c1}) = (8.7 \pm 0.4) \times 10^{-2}$ .			

$$\Gamma(\chi_{c1}(1P) \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{17}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>19.7 ± 1.6 OUR FIT</b>			
<b>13.2 ± 2.4 ± 3.2</b>	<sup>1</sup> BAI	99B	BES $\psi(2S) \rightarrow \gamma K_S^0 K^+ \pi^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c1} \rightarrow K_S^0 K^+ \pi^-)$ reported by BAI 99B was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.8)\%$ and $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$ [BAI 98D].			

$$\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- K^+ K^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}} \times \Gamma_{36}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{133}^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.52 ± 0.11 OUR FIT</b>				
<b>0.61 ± 0.11 ± 0.08</b>	54	<sup>1</sup> ABLIKIM	06T	BES2 $\psi(2S) \rightarrow \gamma K^+ K^+ K^- K^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c1} \rightarrow 2K^+ 2K^-)$ reported by ABLIKIM 06T was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.8)\%$ .				

$$\Gamma(\chi_{c1}(1P) \rightarrow K^+ K^- K^+ K^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{36}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$$

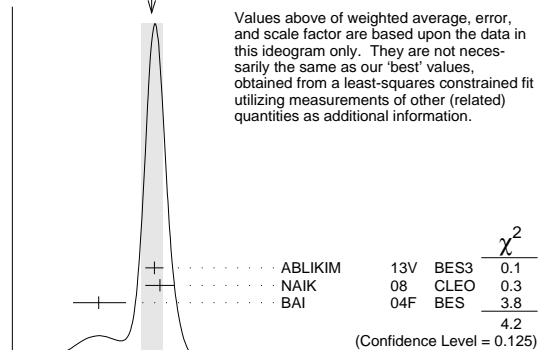
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.52 ± 0.31 OUR FIT</b>			
<b>1.13 ± 0.40 ± 0.29</b>	<sup>1</sup> BAI	99B	BES $\psi(2S) \rightarrow \gamma K^+ K^+ K^- K^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c1} \rightarrow 2K^+ 2K^-)$ reported by BAI 99B was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c1}(1P)) = (8.7 \pm 0.8)\%$ and $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$ [BAI 98D].			

$$\Gamma(\chi_{c1}(1P) \rightarrow \rho \bar{\rho})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}} \times \Gamma_{45}/\Gamma \times \Gamma_{133}^{\psi(2S)}/\Gamma_{133}^{\psi(2S)}$$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.4 ± 0.4 OUR FIT</b>				
<b>7.8 ± 0.6 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
7.9 ± 0.4 ± 0.3	453	ABLIKIM	13v	BES3 $\psi(2S) \rightarrow \gamma \rho \bar{\rho}$
8.2 ± 0.7 ± 0.4	141 ± 13	<sup>1</sup> NAIK	08	CLEO $\psi(2S) \rightarrow \gamma \rho \bar{\rho}$
4.8 ± 1.4 ± 0.6	18.2 <sup>+5.5</sup> <sub>-4.9</sub>	BAI	04F	BES $\psi(2S) \rightarrow \gamma \chi_{c1}(1P) \rightarrow \gamma \rho \bar{\rho}$

<sup>1</sup> Calculated by us. NAIK 08 reports  $B(\chi_{c1} \rightarrow \rho \bar{\rho}) = (9.0 \pm 0.8 \pm 0.4 \pm 0.5) \times 10^{-5}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c1}) = (9.07 \pm 0.11 \pm 0.54)\%$ .

WEIGHTED AVERAGE  
7.8±0.6 (Error scaled by 1.4)



$$\Gamma(\chi_{c1}(1P) \rightarrow \rho \bar{\rho})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c1}(1P))/\Gamma_{\text{total}} \text{ (units } 10^{-6})$$

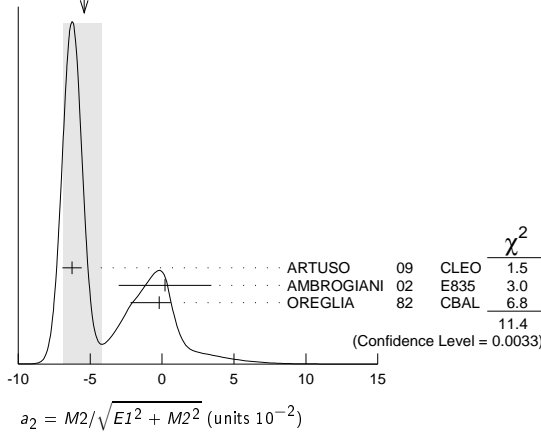
See key on page 601

## Meson Particle Listings

$$\chi_{c1}(1P), h_c(1P)$$

MULTIPOLE AMPLITUDES IN  $\chi_{c1}(1P) \rightarrow \gamma J/\psi(1S)$ 
 $a_2 = M_2/\sqrt{E_1^2 + M_2^2}$  Magnetic quadrupole fractional transition amplitude

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-5.4 \pm 1.2</math></b>	<b>1.5</b>	<b>OUR AVERAGE</b> Error includes scale factor of 2.4. See the ideogram below		
$-6.26 \pm 0.63 \pm 0.24$	39k			
$0.2 \pm 3.2 \pm 0.4$	2090	ARTUSO	09	CLEO $\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
		AMBROGIANI	02	E835 $p\bar{p} \rightarrow \chi_{c1} \rightarrow J/\psi \gamma$
$-0.2 \pm 0.8$	921	OREGLIA	82	CBAL $\psi(2S) \rightarrow \chi_{c1} \gamma \rightarrow J/\psi \gamma \gamma$
$-2.0$				

 WEIGHTED AVERAGE  
 $-5.4 \pm 1.2 - 1.5$  (Error scaled by 2.4)
MULTIPOLE AMPLITUDES IN  $\psi(2S) \rightarrow \gamma \chi_{c1}(1S)$  RADIATIVE DECAY
 $b_2 = M_2/\sqrt{E_1^2 + M_2^2}$  Magnetic quadrupole fractional transition amplitude

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.9 \pm 0.8</math></b>	<b>OUR AVERAGE</b>			
$2.76 \pm 0.73 \pm 0.23$	39k	ARTUSO	09	CLEO $\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
$7.7 \pm 5.0$	921	OREGLIA	82	CBAL $\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
$-4.5$				

MULTIPOLE AMPLITUDE RATIOS IN RADIATIVE DECAYS  
 $\psi(2S) \rightarrow \gamma \chi_{c1}(1S)$  and  $\chi_{c1} \rightarrow \gamma J/\psi(1S)$ 
 $a_2/b_2$  Magnetic quadrupole transition amplitude ratio

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$-2.27^{+0.57}_{-0.99}$	39k	<sup>1</sup> ARTUSO	09	CLEO $\psi(2S) \rightarrow \gamma\gamma\ell^+\ell^-$

<sup>1</sup> Statistical and systematic errors combined. Not independent of  $a_2(\chi_{c1})$  and  $b_2(\chi_{c1})$  values from ARTUSO 09.
 $\chi_{c1}(1P)$  REFERENCES

ABLIKIM	151	PR D91 092006	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15M	PR D91 112008	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	14J	PR D89 074030	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	13B	PR D87 012002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	13D	PR D87 012007	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	13H	PR D87 032007	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	13V	PR D88 112001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12I	PR D86 052004	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12J	PR D86 052011	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12O	PRL 109 172002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11A	PR D83 012006	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11D	PR D83 032003	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11E	PR D83 112005	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11F	PR D83 112009	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	11K	PRL 107 092001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ONYISI	10	PR D82 011103	P.U.E. Onyisi <i>et al.</i>	(CLEO Collab.)
ARTUSO	09	PR D80 112003	M. Artuso <i>et al.</i>	(CLEO Collab.)
BENNETT	08A	PRL 101 151801	J.V. Bennett <i>et al.</i>	(CLEO Collab.)
ECKLUND	08A	PR D78 091501	K.M. Ecklund <i>et al.</i>	(CLEO Collab.)
HE	08B	PR D78 092004	Q. He <i>et al.</i>	(CLEO Collab.)
MENDEZ	08	PR D78 011102	H. Mendez <i>et al.</i>	(CLEO Collab.)
NAIK	08	PR D78 031101	P. Naik <i>et al.</i>	(CLEO Collab.)
ATHAR	07	PR D75 032002	S.B. Athar <i>et al.</i>	(CLEO Collab.)
ABLIKIM	06D	PR D73 052006	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	06R	PR D74 072001	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	06T	PL B642 197	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05G	PR D71 092002	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	05O	PL B630 21	M. Ablikim <i>et al.</i>	(BES Collab.)
ADAM	05A	PRL 94 232002	N.E. Adam <i>et al.</i>	(CLEO Collab.)
ANDREOTTI	05A	NP B717 34	M. Andreotti <i>et al.</i>	(FNAL E835 Collab.)
ABLIKIM	04B	PR D70 012003	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	04H	PR D70 092003	M. Ablikim <i>et al.</i>	(BES Collab.)
ATHAR	04	PR D70 112002	S.B. Athar <i>et al.</i>	(CLEO Collab.)
BAI	04F	PR D69 092001	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	04I	PR D70 012006	J.Z. Bai <i>et al.</i>	(BES Collab.)
AULCHENKO	03	PL B573 63	V.M. Aulchenko <i>et al.</i>	(KEDR Collab.)
BAI	02E	PR D67 112001	J.Z. Bai <i>et al.</i>	(BES Collab.)
AMBROGIANI	02	PR D65 052002	M. Ambrogiani <i>et al.</i>	(FNAL E835 Collab.)
BAI	99B	PR D60 072001	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	98D	PR D58 092006	J.Z. Bai <i>et al.</i>	(BES Collab.)

BAI	98I	PRL 81 3091	J.Z. Bai <i>et al.</i>	(BES Collab.)
ARMSTRONG	92	NP B373 35	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
Also		PRL 68 1468	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
BAGLIN	86B	PL B172 455	C. Baglin	(LAPP, CERN, GENO, LYON, OSLO+)
GAISER	86	PR D34 711	J. Gaiser <i>et al.</i>	(Crystal Ball Collab.)
LEMOIGNE	82	PL 113B 509	Y. Lemoigne <i>et al.</i>	(SACL, LOIC, SHMP+)
OREGLIA	82	PR D25 2259	M.J. Oreglia <i>et al.</i>	(SLAC, CIT, HARV+)
Also		Private Comm.	M.J. Oreglia	(EFI)
HIMEL	80	PRL 44 920	T. Himel <i>et al.</i>	(LBL, SLAC)
Also		Private Comm.	G. Trilling	(LBL, UCB)
BRANDELIK	79B	NP B160 426	R. Brandelik <i>et al.</i>	(DASP Collab.)
BARTL	78B	PL 79B 492	W. Bartel <i>et al.</i>	(DESY, HEIDP)
TANENBAUM	78	PR D17 1731	W.M. Tanenbaum <i>et al.</i>	(SLAC, LBL)
Also		Private Comm.	G. Trilling	(LBL, UCB)
BIDDICK	77	PRL 38 1324	C.J. Biddick <i>et al.</i>	(UCSD, UMD, PAVI+)
FELDMAN	77	PRPL 33C 285	G.J. Feldman, M.L. Perl	(LBL, SLAC)
YAMADA	77	Hamburg Conf. 69	S. Yamada	(DASP Collab.)
WHITAKER	76	PRL 37 1596	J.S. Whitaker <i>et al.</i>	(SLAC, LBL)
TANENBAUM	75	PRL 35 1323	W.M. Tanenbaum <i>et al.</i>	(LBL, SLAC)

$$h_c(1P)$$

$$I^G(J^{PC}) = ?^?(1+ -)$$

 Quantum numbers are quark model prediction,  $C = -$  established by  $\eta_c \gamma$  decay.
 $h_c(1P)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3525.38 \pm 0.11</math></b>	<b>OUR AVERAGE</b>			
$3525.31 \pm 0.11 \pm 0.14$	832	<sup>1</sup> ABLIKIM	12N BES3	$\psi(2S) \rightarrow \pi^0 \gamma$ hadrons
$3525.40 \pm 0.13 \pm 0.18$	3679	ABLIKIM	10B BES3	$\psi(2S) \rightarrow \pi^0 \gamma \eta_c$
$3525.20 \pm 0.18 \pm 0.12$	1282	<sup>2</sup> DOBBS	08A CLEO	$\psi(2S) \rightarrow \pi^0 \eta_c \gamma$
$3525.8 \pm 0.2 \pm 0.2$	13	ANDREOTTI	05B E835	$\bar{p}p \rightarrow \eta_c \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3525.6 \pm 0.5$	$92 \pm 23$	ADAMS	09 CLEO	$\psi(2S) \rightarrow 2(\pi^+ \pi^- \pi^0)$
$3524.4 \pm 0.6 \pm 0.4$	$168 \pm 40$	<sup>3</sup> ROSNER	05 CLEO	$\psi(2S) \rightarrow \pi^0 \eta_c \gamma$
$3527 \pm 8$	42	ANTONIAZZI	94 E705	$300 \pi^\pm, p \text{ Li} \rightarrow J/\psi \pi^0 X$
$3526.28 \pm 0.18 \pm 0.19$	59	<sup>4</sup> ARMSTRONG	92D E760	$\bar{p}p \rightarrow J/\psi \pi^0$
$3525.4 \pm 0.8 \pm 0.4$	5	BAGLIN	86 SPEC	$\bar{p}p \rightarrow J/\psi X$

<sup>1</sup> With floating width.

<sup>2</sup> Combination of exclusive and inclusive analyses for the reaction  $\psi(2S) \rightarrow \pi^0 h_c \rightarrow \pi^0 \eta_c \gamma$ . This result is the average of DOBBS 08A and ROSNER 05.

<sup>3</sup> Superseded by DOBBS 08A.

<sup>4</sup> Mass central value and systematic error recalculated by us according to Eq. (16) in ARMSTRONG 93B, using the value for the  $\psi(2S)$  mass from AULCHENKO 03.
 $h_c(1P)$  WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.70 \pm 0.28 \pm 0.22</math></b>		832	<sup>5</sup> ABLIKIM	12N BES3	$\psi(2S) \rightarrow \pi^0 \gamma$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 1.44$	90	3679	<sup>6</sup> ABLIKIM	10B BES3	$\psi(2S) \rightarrow \pi^0 \gamma \eta_c$
$< 1$		13	ANDREOTTI	05B E835	$\bar{p}p \rightarrow \eta_c \gamma$
$< 1.1$	90	59	ARMSTRONG	92D E760	$\bar{p}p \rightarrow J/\psi \pi^0$

<sup>5</sup> With floating mass.

<sup>6</sup> The central value is  $\Gamma = 0.73 \pm 0.45 \pm 0.28$  MeV.
 $h_c(1P)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $J/\psi(1S) \pi^0$		
$\Gamma_2$ $J/\psi(1S) \pi \pi$	not seen	
$\Gamma_3$ $p\bar{p}$	$< 1.5 \times 10^{-4}$	90%
$\Gamma_4$ $\eta_c(1S) \gamma$	$(51 \pm 6) \%$	
$\Gamma_5$ $\pi^+ \pi^- \pi^0$	$< 2.2 \times 10^{-3}$	
$\Gamma_6$ $2\pi^+ 2\pi^- \pi^0$	$(2.2 \pm 0.8) \%$	
$\Gamma_7$ $3\pi^+ 3\pi^- \pi^0$	$< 2.9 \%$	

 $h_c(1P)$  PARTIAL WIDTHS $h_c(1P) \Gamma(i)\Gamma(\bar{p}p)/\Gamma(\text{total})$ 

$\Gamma(\eta_c(1S)\gamma) \times \Gamma(p\bar{p})/\Gamma_{\text{total}}$				$\Gamma_4\Gamma_3/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$12.0 \pm 4.5$	13	<sup>7</sup> ANDREOTTI	05B E835	$\bar{p}p \rightarrow \eta_c \gamma$
<sup>7</sup> Assuming $\Gamma = 1$ MeV.				

 $h_c(1P)$  BRANCHING RATIOS

$\Gamma(J/\psi(1S)\pi\pi)/\Gamma(J/\psi(1S)\pi^0)$					$\Gamma_2/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.18</b>	90	ARMSTRONG 92D	E760	$\bar{p}p \rightarrow J/\psi \pi^0$	

# Meson Particle Listings

## $h_c(1P)$ , $\chi_{c2}(1P)$

$\Gamma(h_c(1S)\gamma)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	

**51 ± 6 OUR AVERAGE**

54.3± 6.7±5.2 3679 8 ABLIKIM 10B BES3  $\psi(2S) \rightarrow \pi^0 \gamma \eta_c$

48 ± 6 ± 7 8 DOBBS 08A CLEO  $\psi(2S) \rightarrow \pi^0 \eta_c \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

48 ± 6 ± 7 1282 9 DOBBS 08A CLEO  $\psi(2S) \rightarrow \pi^0 \eta_c \gamma$

46 ± 12 ± 7 168 10 ROSNER 05 CLEO  $\psi(2S) \rightarrow \pi^0 \eta_c \gamma$

<sup>8</sup> Average of DOBBS 08A and ROSNER 05. DOBBS 08A reports  $[\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))]$  =  $(4.16 \pm 0.30 \pm 0.37) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$  =  $(8.6 \pm 1.3) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>9</sup> DOBBS 08A reports  $[\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))]$  =  $(4.19 \pm 0.32 \pm 0.45) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$  =  $(8.6 \pm 1.3) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>10</sup> ROSNER 05 reports  $[\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))]$  =  $(4.0 \pm 0.8 \pm 0.7) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$  =  $(8.6 \pm 1.3) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		

**<2.2**

<sup>11</sup> ADAMS 09 CLEO  $\psi(2S) \rightarrow \pi^0 \gamma \eta_c$

<sup>11</sup> ADAMS 09 reports  $[\Gamma(h_c(1P) \rightarrow \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))]$  <  $0.19 \times 10^{-5}$  which we divide by our best value  $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$  =  $8.6 \times 10^{-4}$ .

$\Gamma(2\pi^+ 2\pi^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	

**2.2 ± 0.8 ± 0.3**

92 <sup>12</sup> ADAMS 09 CLEO  $\psi(2S) \rightarrow \pi^0 \gamma \eta_c$

<sup>12</sup> ADAMS 09 reports  $[\Gamma(h_c(1P) \rightarrow 2\pi^+ 2\pi^- \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))]$  =  $(1.88^{+0.48+0.47}_{-0.45-0.30}) \times 10^{-5}$  which we divide by our best value  $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$  =  $(8.6 \pm 1.3) \times 10^{-4}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+ 3\pi^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT		

**<2.9**

<sup>13</sup> ADAMS 09 CLEO  $\psi(2S) \rightarrow \pi^0 \gamma \eta_c$

<sup>13</sup> ADAMS 09 reports  $[\Gamma(h_c(1P) \rightarrow 3\pi^+ 3\pi^- \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \pi^0 h_c(1P))]$  <  $2.5 \times 10^{-5}$  which we divide by our best value  $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$  =  $8.6 \times 10^{-4}$ .

$\Gamma(h_c(1P) \rightarrow \eta_c(1S)\gamma)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma \times \Gamma_{15}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	

**4.3 ± 0.4 OUR AVERAGE**

4.58±0.40±0.50 3679 <sup>14</sup> ABLIKIM 10B BES3  $\psi(2S) \rightarrow \pi^0 \gamma X$

4.16±0.30±0.37 1430 <sup>15</sup> DOBBS 08A CLEO  $\psi(2S) \rightarrow \pi^0 \gamma \eta_c$

<sup>14</sup> Not independent of other branching fractions in ABLIKIM 10B.

<sup>15</sup> Not independent of other branching fractions in DOBBS 08A.

$\Gamma(h_c(1P) \rightarrow \rho \bar{\rho})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \pi^0 h_c(1P))/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma \times \Gamma_{15}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	

**<1.3 × 10<sup>-7</sup>**

90 ABLIKIM 13V BES3  $\psi(2S) \rightarrow \gamma \rho \bar{\rho}$

### $h_c(1P)$ REFERENCES

ABLIKIM	13V	PR D88 112001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12N	PR D86 092009	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	10B	PRL 104 132002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ADAMS	09	PR D80 051106	G.S. Adams <i>et al.</i>	(CLEO Collab.)
DOBBS	08A	PRL 101 182003	S. Dobbs <i>et al.</i>	(CLEO Collab.)
ANDREOTTI	05B	PR D72 032001	M. Andreotti <i>et al.</i>	(FNAL E835 Collab.)
ROSNER	05	PRL 95 102003	J.L. Rosner <i>et al.</i>	(CLEO Collab.)
AULCHENKO	03	PL B573 63	V.M. Aulchenko <i>et al.</i>	(KEDR Collab.)
ANTONIAZZI	94	PR D50 4258	L. Antoniazzi <i>et al.</i>	(E705 Collab.)
ARMSTRONG	93B	PR D47 772	T.A. Armstrong <i>et al.</i>	(FNAL E760 Collab.)
ARMSTRONG	92D	PRL 69 2337	T.A. Armstrong <i>et al.</i>	(FNAL FERR, GENO+)
BAGLIN	86	PL B171 135	C. Baglin <i>et al.</i>	(LAPP, CERN, TORI, STRB+)

## $\chi_{c2}(1P)$

See the Review on " $\psi(2S)$  and  $\chi_c$  branching ratios" before the  $\chi_{c0}(1P)$  Listings.

$$I^G(J^{PC}) = 0^+(2^{++})$$

### $\chi_{c2}(1P)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3556.20 ± 0.09 OUR AVERAGE</b>				
3555.3 ± 0.6 ± 2.2	25k	UEHARA	08	BELL $\gamma\gamma \rightarrow$ hadrons
3555.70 ± 0.59 ± 0.39		ABLIKIM	05G	BES2 $\psi(2S) \rightarrow \gamma \chi_{c2}$
3556.173 ± 0.123 ± 0.020		ANDREOTTI	05A	E835 $p\bar{p} \rightarrow e^+ e^- \gamma$
3559.9 ± 2.9		EISENSTEIN	01	CLE2 $e^+ e^- \rightarrow e^+ e^- \chi_{c2}$
3556.4 ± 0.7		BAI	99B	BES $\psi(2S) \rightarrow \gamma X$
3556.22 ± 0.131 ± 0.020	585	<sup>1</sup> ARMSTRONG	92	E760 $\bar{p}p \rightarrow e^+ e^- \gamma$
3556.9 ± 0.4 ± 0.5	50	BAGLIN	86B	SPEC $\bar{p}p \rightarrow e^+ e^- X$
3557.8 ± 0.2 ± 4		<sup>2</sup> GAISER	86	CBAL $\psi(2S) \rightarrow \gamma X$
3553.4 ± 2.2	66	<sup>3</sup> LEMOIGNE	82	GOLI $185 \pi^- \text{Be} \rightarrow \gamma \mu^+ \mu^- \text{A}$
3555.9 ± 0.7		<sup>4</sup> OREGLIA	82	CBAL $e^+ e^- \rightarrow J/\psi 2\gamma$
3557 ± 1.5	69	<sup>5</sup> HIMEL	80	MRK2 $e^+ e^- \rightarrow J/\psi 2\gamma$
3551 ± 11	15	BRANDELIK	79B	DASP $e^+ e^- \rightarrow J/\psi 2\gamma$
3553 ± 4		<sup>5</sup> BARTEL	78B	CNTR $e^+ e^- \rightarrow J/\psi 2\gamma$
3553 ± 4 ± 4		<sup>5,6</sup> TANENBAUM	78	MRK1 $e^+ e^-$
3563 ± 7	360	<sup>5</sup> BIDDICK	77	CNTR $e^+ e^- \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3555.4 ± 1.3	53	UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$
3543 ± 10	4	WHITAKER	76	MRK1 $e^+ e^- \rightarrow J/\psi 2\gamma$

<sup>1</sup> Recalculated by ANDREOTTI 05A, using the value of  $\psi(2S)$  mass from AULCHENKO 03.

<sup>2</sup> Using mass of  $\psi(2S)$  = 3686.0 MeV.

<sup>3</sup>  $J/\psi(1S)$  mass constrained to 3097 MeV.

<sup>4</sup> Assuming  $\psi(2S)$  mass = 3686 MeV and  $J/\psi(1S)$  mass = 3097 MeV.

<sup>5</sup> Mass value shifted by us by amount appropriate for  $\psi(2S)$  mass = 3686 MeV and  $J/\psi(1S)$  mass = 3097 MeV.

<sup>6</sup> From a simultaneous fit to radiative and hadronic decay channels.

### $\chi_{c2}(1P)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.93 ± 0.11 OUR FIT</b>				
<b>1.95 ± 0.13 OUR AVERAGE</b>				
1.915 ± 0.188 ± 0.013		ANDREOTTI	05A	E835 $p\bar{p} \rightarrow e^+ e^- \gamma$
1.96 ± 0.17 ± 0.07	585	<sup>1</sup> ARMSTRONG	92	E760 $\bar{p}p \rightarrow e^+ e^- \gamma$
2.6 <sup>+1.4</sup> <sub>-1.0</sub>	50	BAGLIN	86B	SPEC $\bar{p}p \rightarrow e^+ e^- X$
2.8 <sup>+2.1</sup> <sub>-2.0</sub>		<sup>2</sup> GAISER	86	CBAL $\psi(2S) \rightarrow \gamma X$

<sup>1</sup> Recalculated by ANDREOTTI 05A.

<sup>2</sup> Errors correspond to 90% confidence level; authors give only width range.

### $\chi_{c2}(1P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
Hadronic decays		
$\Gamma_1$ $2(\pi^+ \pi^-)$	( 1.07 ± 0.10 ) %	
$\Gamma_2$ $\rho \rho$		
$\Gamma_3$ $\pi^+ \pi^- \pi^0 \pi^0$	( 1.91 ± 0.25 ) %	
$\Gamma_4$ $\rho^+ \pi^- \pi^0 + \text{c.c.}$	( 2.3 ± 0.4 ) %	
$\Gamma_5$ $4\pi^0$	( 1.16 ± 0.16 ) × 10 <sup>-3</sup>	
$\Gamma_6$ $K^+ K^- \pi^0 \pi^0$	( 2.2 ± 0.4 ) × 10 <sup>-3</sup>	
$\Gamma_7$ $K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.}$	( 1.44 ± 0.21 ) %	
$\Gamma_8$ $\rho^- K^+ \bar{K}^0 + \text{c.c.}$	( 4.3 ± 1.3 ) × 10 <sup>-3</sup>	
$\Gamma_9$ $K^*(892)^0 K^- \pi^+ \rightarrow K^- \pi^+ + K^0 \pi^0 + \text{c.c.}$	( 3.1 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{10}$ $K^*(892)^0 \bar{K}^0 \pi^- \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.}$	( 4.0 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{11}$ $K^*(892)^- K^+ \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.}$	( 3.9 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{12}$ $K^*(892)^+ \bar{K}^0 \pi^- \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.}$	( 3.1 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{13}$ $K^+ K^- \eta \pi^0$	( 1.3 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{14}$ $K^+ K^- \pi^+ \pi^-$	( 8.9 ± 1.0 ) × 10 <sup>-3</sup>	
$\Gamma_{15}$ $K^+ K^- \pi^+ \pi^- \pi^0$	( 1.17 ± 0.13 ) %	
$\Gamma_{16}$ $K_S^0 K^\pm \pi^\mp \pi^+ \pi^-$	( 7.3 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{17}$ $K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.}$	( 2.2 ± 1.1 ) × 10 <sup>-3</sup>	
$\Gamma_{18}$ $K^*(892)^0 \bar{K}^*(892)^0$	( 2.4 ± 0.5 ) × 10 <sup>-3</sup>	
$\Gamma_{19}$ $3(\pi^+ \pi^-)$	( 8.6 ± 1.8 ) × 10 <sup>-3</sup>	
$\Gamma_{20}$ $\phi \phi$	( 1.12 ± 0.10 ) × 10 <sup>-3</sup>	
$\Gamma_{21}$ $\omega \omega$	( 8.8 ± 1.1 ) × 10 <sup>-4</sup>	



$\Gamma_{22}$	$\omega K^+ K^-$	$(7.3 \pm 0.9) \times 10^{-4}$	
$\Gamma_{23}$	$\omega \phi$		
$\Gamma_{24}$	$\pi \pi$	$(2.33 \pm 0.12) \times 10^{-3}$	
$\Gamma_{25}$	$\rho^0 \pi^+ \pi^-$	$(3.8 \pm 1.6) \times 10^{-3}$	
$\Gamma_{26}$	$\pi^+ \pi^- \eta$	$(5.0 \pm 1.3) \times 10^{-4}$	
$\Gamma_{27}$	$\pi^+ \pi^- \eta'$	$(5.2 \pm 1.9) \times 10^{-4}$	
$\Gamma_{28}$	$\eta \eta$	$(5.7 \pm 0.5) \times 10^{-4}$	
$\Gamma_{29}$	$K^+ K^-$	$(1.05 \pm 0.07) \times 10^{-3}$	
$\Gamma_{30}$	$K_S^0 K_S^0$	$(5.5 \pm 0.4) \times 10^{-4}$	
$\Gamma_{31}$	$\bar{K}^0 K^+ \pi^- + \text{c.c.}$	$(1.34 \pm 0.19) \times 10^{-3}$	
$\Gamma_{32}$	$K^+ K^- \pi^0$	$(3.2 \pm 0.8) \times 10^{-4}$	
$\Gamma_{33}$	$K^+ K^- \eta$	$< 3.4 \times 10^{-4}$	90%
$\Gamma_{34}$	$K^+ K^- \eta'(958)$	$(1.94 \pm 0.34) \times 10^{-4}$	
$\Gamma_{35}$	$\eta \eta'$	$< 6 \times 10^{-5}$	90%
$\Gamma_{36}$	$\eta' \eta'$	$< 1.0 \times 10^{-4}$	90%
$\Gamma_{37}$	$\pi^+ \pi^- K_S^0 K_S^0$	$(2.3 \pm 0.6) \times 10^{-3}$	
$\Gamma_{38}$	$K^+ K^- K_S^0 K_S^0$	$< 4 \times 10^{-4}$	90%
$\Gamma_{39}$	$K^+ K^- K^+ K^-$	$(1.73 \pm 0.21) \times 10^{-3}$	
$\Gamma_{40}$	$K^+ K^- \phi$	$(1.48 \pm 0.31) \times 10^{-3}$	
$\Gamma_{41}$	$\bar{K}^0 K^+ \pi^- \phi + \text{c.c.}$	$(4.8 \pm 0.7) \times 10^{-3}$	
$\Gamma_{42}$	$K^+ K^- \pi^0 \phi$	$(2.7 \pm 0.5) \times 10^{-3}$	
$\Gamma_{43}$	$\phi \pi^+ \pi^- \pi^0$	$(9.3 \pm 1.2) \times 10^{-4}$	
$\Gamma_{44}$	$\rho \bar{\rho}$	$(7.5 \pm 0.4) \times 10^{-5}$	
$\Gamma_{45}$	$\rho \bar{\rho} \pi^0$	$(4.9 \pm 0.4) \times 10^{-4}$	
$\Gamma_{46}$	$\rho \bar{\rho} \eta$	$(1.82 \pm 0.26) \times 10^{-4}$	
$\Gamma_{47}$	$\rho \bar{\rho} \omega$	$(3.8 \pm 0.5) \times 10^{-4}$	
$\Gamma_{48}$	$\rho \bar{\rho} \phi$	$(2.9 \pm 0.9) \times 10^{-5}$	
$\Gamma_{49}$	$\rho \bar{\rho} \pi^+ \pi^-$	$(1.32 \pm 0.34) \times 10^{-3}$	
$\Gamma_{50}$	$\rho \bar{\rho} \pi^0 \pi^0$	$(8.2 \pm 2.5) \times 10^{-4}$	
$\Gamma_{51}$	$\rho \bar{\rho} K^+ K^- (\text{non-resonant})$	$(2.00 \pm 0.34) \times 10^{-4}$	
$\Gamma_{52}$	$\rho \bar{\rho} K_S^0 K_S^0$	$< 7.9 \times 10^{-4}$	90%
$\Gamma_{53}$	$\rho \bar{\rho} \pi^-$	$(8.9 \pm 1.0) \times 10^{-4}$	
$\Gamma_{54}$	$\bar{\rho} \rho \pi^+$	$(9.3 \pm 0.9) \times 10^{-4}$	
$\Gamma_{55}$	$\bar{\rho} \rho \pi^- \pi^0$	$(2.27 \pm 0.19) \times 10^{-3}$	
$\Gamma_{56}$	$\bar{\rho} \rho \pi^+ \pi^0$	$(2.21 \pm 0.20) \times 10^{-3}$	
$\Gamma_{57}$	$\Lambda \bar{\Lambda}$	$(1.92 \pm 0.16) \times 10^{-4}$	
$\Gamma_{58}$	$\Lambda \bar{\Lambda} \pi^+ \pi^-$	$(1.31 \pm 0.17) \times 10^{-3}$	
$\Gamma_{59}$	$\Lambda \bar{\Lambda} \pi^+ \pi^- (\text{non-resonant})$	$(6.9 \pm 1.6) \times 10^{-4}$	
$\Gamma_{60}$	$\Sigma(1385)^+ \bar{\Lambda} \pi^- + \text{c.c.}$	$< 4 \times 10^{-4}$	90%
$\Gamma_{61}$	$\Sigma(1385)^- \bar{\Lambda} \pi^+ + \text{c.c.}$	$< 6 \times 10^{-4}$	90%
$\Gamma_{62}$	$K^+ \bar{\rho} \Lambda + \text{c.c.}$	$(8.1 \pm 0.6) \times 10^{-4}$	
$\Gamma_{63}$	$K^+ \bar{\rho} \Lambda(1520) + \text{c.c.}$	$(2.9 \pm 0.7) \times 10^{-4}$	
$\Gamma_{64}$	$\Lambda(1520) \bar{\Lambda}(1520)$	$(4.8 \pm 1.5) \times 10^{-4}$	
$\Gamma_{65}$	$\Sigma^0 \bar{\Sigma}^0$	$< 6 \times 10^{-5}$	90%
$\Gamma_{66}$	$\Sigma^+ \bar{\Sigma}^-$	$< 7 \times 10^{-5}$	90%
$\Gamma_{67}$	$\Sigma(1385)^+ \bar{\Sigma}(1385)^-$	$< 1.6 \times 10^{-4}$	90%
$\Gamma_{68}$	$\Sigma(1385)^- \bar{\Sigma}(1385)^+$	$< 8 \times 10^{-5}$	90%
$\Gamma_{69}$	$K^- \Lambda \Xi^+ + \text{c.c.}$	$(1.84 \pm 0.34) \times 10^{-4}$	
$\Gamma_{70}$	$\Xi^0 \Xi^0$	$< 1.1 \times 10^{-4}$	90%
$\Gamma_{71}$	$\Xi^- \Xi^+$	$(1.48 \pm 0.33) \times 10^{-4}$	
$\Gamma_{72}$	$J/\psi(1S) \pi^+ \pi^- \pi^0$	$< 1.5 \%$	90%
$\Gamma_{73}$	$\pi^0 \eta_c$	$< 3.2 \times 10^{-3}$	90%
$\Gamma_{74}$	$\eta_c(1S) \pi^+ \pi^-$	$< 5.4 \times 10^{-3}$	90%

### Radiative decays

$\Gamma_{75}$	$\gamma J/\psi(1S)$	$(19.2 \pm 0.7) \%$	
$\Gamma_{76}$	$\gamma \rho^0$	$< 2.0 \times 10^{-5}$	90%
$\Gamma_{77}$	$\gamma \omega$	$< 6 \times 10^{-6}$	90%
$\Gamma_{78}$	$\gamma \phi$	$< 8 \times 10^{-6}$	90%
$\Gamma_{79}$	$\gamma \gamma$	$(2.74 \pm 0.14) \times 10^{-4}$	

### CONSTRAINED FIT INFORMATION

A multiparticle fit to  $\chi_{C1}(1P)$ ,  $\chi_{C0}(1P)$ ,  $\chi_{C2}(1P)$ , and  $\psi(2S)$  with 4 total widths, a partial width, 25 combinations of partial widths obtained from integrated cross section, and 84 branching ratios uses 240 measurements to determine 49 parameters. The overall fit has a  $\chi^2 = 342.4$  for 191 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \cdot \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ .

$x_{14}$	13									
$x_{17}$	3	21								
$x_{18}$	8	7	1							
$x_{20}$	14	12	3	7						
$x_{24}$	19	16	3	10	24					
$x_{25}$	19	3	1	2	3	4				
$x_{28}$	11	9	2	6	14	27	2			
$x_{29}$	14	12	3	7	17	33	3	19		
$x_{30}$	13	11	2	6	15	28	3	17	20	
$x_{31}$	7	6	1	4	8	16	1	9	11	10
$x_{39}$	9	8	2	5	10	18	2	10	13	11
$x_{44}$	16	13	3	8	16	24	4	14	17	15
$x_{57}$	11	9	2	6	14	28	2	16	20	17
$x_{75}$	24	21	4	12	29	55	5	32	40	34
$x_{79}$	-8	-6	-1	-3	1	19	-2	13	13	10
$\Gamma$	-28	-23	-5	-14	-28	-43	-6	-25	-32	-28
	$x_1$	$x_{14}$	$x_{17}$	$x_{18}$	$x_{20}$	$x_{24}$	$x_{25}$	$x_{28}$	$x_{29}$	$x_{30}$

### $\chi_{c2}(1P)$ PARTIAL WIDTHS

$$= \chi_{c2}(1P) \Gamma(i) \Gamma(\gamma J/\psi(1S)) / \Gamma(\text{total})$$

$$\Gamma(p\bar{p}) \times \Gamma(\gamma J/\psi(1S))/\Gamma_{\text{total}} \qquad \Gamma_{44}\Gamma_{75}/\Gamma$$

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b>27.9±1.3 OUR FIT</b>			
<b>27.5±1.5 OUR AVERAGE</b>			
27.0±1.5±1.1	<sup>1</sup> ANDREOTTI	05A E835	$p\bar{p} \rightarrow e^+e^-\gamma$
27.7±1.5±2.0	<sup>1,2</sup> ARMSTRONG	92 E760	$\bar{p}p \rightarrow e^+e^-\gamma$
36 ± 8	<sup>1</sup> BAGLIN	86B SPEC	$\bar{p}p \rightarrow e^+e^-X$

<sup>1</sup> Calculated by us using  $B(J/\psi(1S) \rightarrow e^+ e^-) = 0.0593 \pm 0.0010$ .

<sup>2</sup> Recalculated by ANDREOTTI 05A.

$$\Gamma(\gamma\gamma) \times \Gamma(\gamma J/\psi(1S))/\Gamma_{\text{total}} \qquad \Gamma_{79}\Gamma_{75}/\Gamma$$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>102± 5 OUR FIT</b>				
<b>117± 10 OUR AVERAGE</b>				
111± 12± 9	147 ± 15	<sup>1</sup> DOBBS	06 CLE3	$10.4 e^+ e^- \rightarrow e^+ e^- \chi_{c2}$
114± 11± 9	136 ± 13.3	<sup>1,2</sup> ABE	02T BELL	$e^+ e^- \rightarrow e^+ e^- \chi_{c2}$
139± 55± 21		<sup>1,3</sup> ACCIARRI	99E L3	$e^+ e^- \rightarrow e^+ e^- \chi_{c2}$
242± 65± 51		<sup>1,4</sup> ACKER...,K...	98 OPAL	$e^+ e^- \rightarrow e^+ e^- \chi_{c2}$
150± 42± 36		<sup>1,5</sup> DOMINICK	94 CLE2	$e^+ e^- \rightarrow e^+ e^- \chi_{c2}$
470± 240± 120		<sup>1,6</sup> BAUER	93 TPC	$e^+ e^- \rightarrow e^+ e^- \chi_{c2}$

<sup>1</sup> Calculated by us using  $B(J/\psi \rightarrow \ell^+ \ell^-) = 0.1187 \pm 0.0008$ .

<sup>2</sup> All systematic errors added in quadrature.

<sup>3</sup>The value for  $\Gamma(\chi_{C2} \rightarrow \gamma\gamma)$  reported in ACCIARRI 99E is derived using  $B(\chi_{C2} \rightarrow \gamma J/\psi(1S)) \times B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.0162 \pm 0.0014$ .

<sup>4</sup> The value for  $\Gamma(\chi_{C2} \rightarrow \gamma\gamma)$  reported in ACKERSTAFF, K 98 is derived using  $B(\chi_{C2} \rightarrow \gamma J/\psi(1S)) = 0.135 \pm 0.011$  and  $B(J/\psi(1S) \rightarrow \ell^+\ell^-) = 0.1203 \pm 0.0038$ .

5 The value for  $\Gamma(\chi_{C2} \rightarrow \gamma\gamma)$  reported in DOMINICK 94 is derived using  $B(\chi_{C2} \rightarrow \gamma J/\psi(1S)) = 0.135 \pm 0.011$ ,  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0627 \pm 0.0020$ , and  $B(J/\psi(1S) \rightarrow \mu^+\mu^-) = 0.0597 \pm 0.0025$ .

<sup>6</sup> The value for  $\Gamma(\chi_{c2} \rightarrow \gamma\gamma)$  reported in BAUER 93 is derived using  $B(\chi_{c2} \rightarrow \gamma J/\psi(1S)) = 0.135 \pm 0.011$ ,  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0627 \pm 0.0020$ , and  $B(J/\psi(1S) \rightarrow \mu^+\mu^-) = 0.0597 \pm 0.0025$ .

$$- \chi_{c2}(1P) \Gamma(i) \Gamma(\gamma\gamma) / \Gamma(\text{total})$$

$$\Gamma(2(\pi^+\pi^-)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}} \qquad \Gamma_1\Gamma_{79}/\Gamma$$

VALUE (eV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>5.7 ± 0.5 OUR FIT</b>				
<b>5.2 ± 0.7 OUR AVERAGE</b>				
5.01 ± 0.44 ± 0.55	1597 ± 138	UEHARA	08 BELL	$\gamma\gamma \rightarrow \chi_{C2} \rightarrow 2(\pi^+\pi^-)$
6.4 ± 1.8 ± 0.8		EISENSTEIN	01 CLE2	$e^+e^- \rightarrow e^+e^-\chi_{C2}$

$$\Gamma(\rho\rho) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}} \qquad \Gamma_2\Gamma_{79}/\Gamma$$

VALUE (eV)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
<7.8	90	<598	UEHARA	08	BELL $\gamma\gamma \rightarrow \chi_{c2} \rightarrow 2(\pi^+\pi^-)$

# Meson Particle Listings

## $\chi_{c2}(1P)$

$\Gamma(K^+K^-\pi^+\pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{14}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.7 ± 0.5 OUR FIT</b>					
<b>4.42±0.42±0.53</b>	780 ± 74	UEHARA	08	BELL $\gamma\gamma \rightarrow \chi_{c2} \rightarrow K^+K^-\pi^+\pi^-$	

$\Gamma(K^+K^-\pi^+\pi^0) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{15}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>6.5±0.9±1.5</b>	1250	DEL-AMO-SA...11M	BABR	$\gamma\gamma \rightarrow K^+K^-\pi^+\pi^0$	

$\Gamma(K^*(892)^0\bar{K}^*(892)^0) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{18}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.26±0.24 OUR FIT</b>					
<b>0.8 ± 0.17±0.27</b>	151 ± 30	UEHARA	08	BELL $\gamma\gamma \rightarrow \chi_{c2} \rightarrow K^+K^-\pi^+\pi^-$	

$\Gamma(\phi\phi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{20}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.59±0.05 OUR FIT</b>					
<b>0.62±0.07±0.05</b>	89 ± 11	<sup>1</sup> LIU	12B	BELL $\gamma\gamma \rightarrow 2(K^+K^-)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.58±0.18±0.16	26.5 ± 8.1	UEHARA	08	BELL $\gamma\gamma \rightarrow \chi_{c2} \rightarrow 2(K^+K^-)$	
<sup>1</sup> Supersedes UEHARA 08. Using $B(\phi \rightarrow K^+K^-) = (48.9 \pm 0.5)\%$ .					

$\Gamma(\omega\omega) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{21}\Gamma_{79}/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.64	90	<sup>1</sup> LIU	12B	BELL $\gamma\gamma \rightarrow 2(\pi^+\pi^-\pi^0)$	
<sup>1</sup> Using $B(\omega \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7)\%$ .					

$\Gamma(\omega\phi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{23}\Gamma_{79}/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.04	90	<sup>1</sup> LIU	12B	BELL $\gamma\gamma \rightarrow K^+K^-\pi^+\pi^-\pi^0$	
<sup>1</sup> Using $B(\phi \rightarrow K^+K^-) = (48.9 \pm 0.5)\%$ and $B(\omega \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7)\%$ .					

$\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{24}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.23±0.08 OUR FIT</b>					
<b>1.18±0.25 OUR AVERAGE</b>					
1.44±0.54±0.47	34 ± 13	<sup>1</sup> UEHARA	09	BELL $10.6\,e^+e^- \rightarrow e^+e^-\pi^0\pi^0$	
1.14±0.21±0.17	54 ± 10	<sup>2</sup> NAKAZAWA	05	BELL $10.6\,e^+e^- \rightarrow e^+e^-\pi^+\pi^-$	
<sup>1</sup> We multiplied the measurement by 3 to convert from $\pi^0\pi^0$ to $\pi\pi$ . Interference with the continuum included.					
<sup>2</sup> We have multiplied $\pi^+\pi^-$ measurement by 3/2 to obtain $\pi\pi$ .					

$\Gamma(\rho^0\pi^+\pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{25}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.0±0.9 OUR FIT</b>					
<b>3.2±1.9±0.5</b>	986 ± 578	UEHARA	08	BELL $\gamma\gamma \rightarrow \chi_{c2} \rightarrow 2(\pi^+\pi^-)$	

$\Gamma(\eta\eta) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{28}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.53±0.22±0.09</b>	8	<sup>1</sup> UEHARA	10A	BELL $10.6\,e^+e^- \rightarrow e^+e^-\eta\eta$	
<sup>1</sup> Interference with the continuum not included.					

$\Gamma(K^+K^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{29}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.56±0.04 OUR FIT</b>					
<b>0.44±0.11±0.07</b>	33 ± 8	NAKAZAWA	05	BELL $10.6\,e^+e^- \rightarrow e^+e^-K^+K^-$	

$\Gamma(K_S^0K_S^0) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{30}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.291±0.025 OUR FIT</b>					
<b>0.27 <math>\pm</math> 0.07 <math>\pm</math> 0.03</b>	53	<sup>1</sup> UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0K_S^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.31 ± 0.05 ± 0.03	38 ± 7	CHEN	07B	BELL $e^+e^- \rightarrow e^+e^-\chi_{c2}$	
<sup>1</sup> Supersedes CHEN 07B.					

$\Gamma(\bar{K}^0K^+\pi^- + \text{c.c.}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{31}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.71±0.11 OUR FIT</b>					
<b>1.20±0.33±0.13</b>	126	<sup>1</sup> DEL-AMO-SA...11M	BABR	$\gamma\gamma \rightarrow K_S^0K^\pm\pi^\mp$	
<sup>1</sup> We have multiplied $\bar{K}K\pi$ by 2/3 to obtain $\bar{K}^0K^+\pi^- + \text{c.c.}$					

$\Gamma(K^+K^-K^+K^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{39}\Gamma_{79}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.91±0.12 OUR FIT</b>					
<b>1.10±0.21±0.15</b>	126 ± 24	UEHARA	08	BELL $\gamma\gamma \rightarrow \chi_{c2} \rightarrow 2(K^+K^-)$	

$\Gamma(\eta_{c15})\pi^+\pi^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{74}\Gamma_{79}/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;15.7</b>	90	LEES	12AE	BABR $e^+e^- \rightarrow e^+e^-\pi^+\pi^-\eta_c$	

## $\chi_{c2}(1P)$ BRANCHING RATIOS

### HADRONIC DECAYS

$\Gamma(2(\pi^+\pi^-))/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	
<b>0.0107±0.0010 OUR FIT</b>		

$\Gamma(\rho^0\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$			$\Gamma_{25}/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.36±0.15 OUR FIT</b>			
<b>0.31±0.17</b>	TANENBAUM 78	MRK1	$\psi(2S) \rightarrow \gamma\chi_{c2}$

$\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.91±0.24±0.07</b>	903.5	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$

<sup>1</sup> HE 08B reports  $1.87 \pm 0.07 \pm 0.22 \pm 0.13\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\rho^+\pi^-\pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.28±0.35±0.08</b>	1031.9	<sup>1,2</sup> HE	08B	CLEO $e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $2.23 \pm 0.11 \pm 0.32 \pm 0.16\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \rho^+\pi^-\pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Calculated by us. We have added the values from HE 08B for  $\rho^+\pi^-\pi^0$  and  $\rho^-\pi^+\pi^0$  decays assuming uncorrelated statistical and fully correlated systematic uncertainties.

$\Gamma(4\pi^0)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.16±0.15±0.04</b>	1164	<sup>1</sup> ABLIKIM	11A BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c2}$

<sup>1</sup> ABLIKIM 11A reports  $(1.21 \pm 0.05 \pm 0.16) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow 4\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+K^-\pi^0\pi^0)/\Gamma_{\text{total}}$		$\Gamma_6/\Gamma$		
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22±0.04±0.01</b>	76.9	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$

<sup>1</sup> HE 08B reports  $0.21 \pm 0.03 \pm 0.03 \pm 0.01\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+K^-\pi^0\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.44±0.20±0.05</b>	211.6	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $1.41 \pm 0.11 \pm 0.16 \pm 0.10\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+\pi^-\bar{K}^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\rho^- K^+ \bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.43±0.13±0.01</b>	62.9	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$

<sup>1</sup> HE 08B reports  $0.42 \pm 0.11 \pm 0.06 \pm 0.03\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \rho^-K^+\bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^*(892)^0 K^- \pi^+ \rightarrow K^- \pi^+ K^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.31±0.08±0.01</b>	38.7	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $0.30 \pm 0.07 \pm 0.04 \pm 0.02\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^*(892)^0K^-\pi^+ \rightarrow K^-\pi^+K^0\pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

See key on page 601

# Meson Particle Listings

 $\chi_{c2}(1P)$ 

$\Gamma(K^*(892)^0 \bar{K}^0 \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{10}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.40±0.09±0.01</b>	63.0	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $0.39 \pm 0.07 \pm 0.05 \pm 0.03$  % from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^*(892)^0 \bar{K}^0 \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^*(892)^- K^+ \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.39±0.08±0.01</b>	51.1	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $0.38 \pm 0.07 \pm 0.04 \pm 0.03$  % from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^*(892)^- K^+ \pi^0 \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^*(892)^+ \bar{K}^0 \pi^- \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.31±0.08±0.01</b>	39.3	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $0.30 \pm 0.07 \pm 0.04 \pm 0.02$  % from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^*(892)^+ \bar{K}^0 \pi^- \rightarrow K^+ \pi^- \bar{K}^0 \pi^0 + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+ K^- \eta \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.133±0.046±0.005</b>	22.9	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$	

<sup>1</sup> HE 08B reports  $0.13 \pm 0.04 \pm 0.02 \pm 0.01$  % from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- \eta \pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{14}/\Gamma$			
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b>8.9±1.0 OUR FIT</b>					

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>11.69±0.13±1.31</b>	11k	<sup>1</sup> ABLIKIM	13B BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$	
<sup>1</sup> Using $1.06 \times 10^8$ $\psi(2S)$ mesons and $\text{B}(\psi(2S) \rightarrow \chi_{c2} \gamma) = (8.72 \pm 0.34)\%$ .					

$\Gamma(K_S^0 K_S^\pm \pi^\mp \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{16}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>7.30±0.11±0.75</b>	4.5k	<sup>1</sup> ABLIKIM	13B BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$	
<sup>1</sup> Using $1.06 \times 10^8$ $\psi(2S)$ mesons and $\text{B}(\psi(2S) \rightarrow \chi_{c2} \gamma) = (8.72 \pm 0.34)\%$ .					

$\Gamma(K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.})/\Gamma(K^+ K^- \pi^+ \pi^-)$					$\Gamma_{17}/\Gamma_{14}$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.25±0.13 OUR FIT</b>					
<b>0.25±0.13</b>	TANENBAUM 78	MRK1	$\psi(2S) \rightarrow \gamma \chi_{c2}$		

$\Gamma(K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID				
<b>22±11 OUR FIT</b>					

$\Gamma(K^*(892)^0 \bar{K}^*(892)^0)/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b>2.4±0.5 OUR FIT</b>					

$\Gamma(3(\pi^+ \pi^-))/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b>8.6±1.8 OUR EVALUATION</b>	Treating systematic error as correlated.				
<b>8.6±1.8 OUR AVERAGE</b>					
$8.6 \pm 0.9 \pm 1.6$	<sup>1</sup> BAI	99B BES	$\psi(2S) \rightarrow \gamma \chi_{c2}$		
$8.7 \pm 5.9 \pm 0.4$	<sup>1</sup> TANENBAUM 78	MRK1	$\psi(2S) \rightarrow \gamma \chi_{c2}$		

<sup>1</sup> Rescaled by us using  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}) = (8.3 \pm 0.4)\%$  and  $\text{B}(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (32.6 \pm 0.5)\%$ . Multiplied by a factor of 2 to convert from  $K_S^0 K^+ \pi^-$  to  $K^0 K^+ \pi^-$  decay.

$\Gamma(\phi\phi)/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b>1.12±0.10 OUR FIT</b>					

$\Gamma(\omega\omega)/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.88±0.11 OUR AVERAGE</b>					
$0.85 \pm 0.10 \pm 0.03$	762	<sup>1</sup> ABLIKIM	11K BES3	$\psi(2S) \rightarrow \gamma$ hadrons	
$1.8 \pm 0.6 \pm 0.1$	$27.7 \pm 7.4$	<sup>2</sup> ABLIKIM	05N BES2	$\psi(2S) \rightarrow \gamma \chi_{c2} \rightarrow \gamma 6\pi$	

<sup>1</sup> ABLIKIM 11k reports  $(8.9 \pm 0.3 \pm 1.1) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \omega\omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ABLIKIM 05N reports  $[\Gamma(\chi_{c2}(1P) \rightarrow \omega\omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  =  $(0.165 \pm 0.044 \pm 0.032) \times 10^{-3}$  which we divide by our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\omega K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.73±0.04±0.08</b>	512	<sup>1</sup> ABLIKIM	13B BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$	
<sup>1</sup> Using $1.06 \times 10^8$ $\psi(2S)$ mesons and $\text{B}(\psi(2S) \rightarrow \chi_{c2} \gamma) = (8.72 \pm 0.34)\%$ .					

$\Gamma(\omega\phi)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.9	90	<sup>1</sup> ABLIKIM	11K BES3	$\psi(2S) \rightarrow \gamma$ hadrons	
<sup>1</sup> ABLIKIM 11k reports $< 2 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{c2}(1P) \rightarrow \omega\phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$ assuming $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .					

$\Gamma(\pi\pi)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b>2.33±0.12 OUR FIT</b>					

$\Gamma(\rho^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID				
<b>38±16 OUR FIT</b>					

$\Gamma(\pi^+ \pi^- \eta)/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>0.50±0.13±0.02</b>		<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.5	90	<sup>2</sup> ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \chi_{c2}$	

<sup>1</sup> ATHAR 07 reports  $(0.49 \pm 0.12 \pm 0.06) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \pi^+ \pi^- \eta)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ABLIKIM 06R reports  $< 1.7 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \pi^+ \pi^- \eta)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.1 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\pi^+ \pi^- \eta')/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b>0.52±0.19±0.02</b>	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$		

<sup>1</sup> ATHAR 07 reports  $(0.51 \pm 0.18 \pm 0.06) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \pi^+ \pi^- \eta')/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\eta\eta)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID				
<b>5.7±0.5 OUR FIT</b>					

$\Gamma(K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b>1.05±0.07 OUR FIT</b>					

$\Gamma(K_S^0 K_S^0)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID				
<b>0.55±0.04 OUR FIT</b>					

$\Gamma(K_S^0 K_S^0)/\Gamma(\pi\pi)$					$\Gamma_{30}/\Gamma_{24}$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.235±0.019 OUR FIT</b>					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.27 \pm 0.07 \pm 0.04$	<sup>1,2</sup> CHEN	07B BELL	$e^+e^- \rightarrow e^+e^- \chi_{c2}$		

<sup>1</sup> Using  $\Gamma(\pi\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  from the  $\pi^+ \pi^-$  measurement of NAKAZAWA 05 rescaled by 3/2 to convert to  $\pi\pi$ .

<sup>2</sup> Not independent from other measurements.

# Meson Particle Listings

## $\chi_{c2}(1P)$

### $\Gamma(K_S^0 K_S^0)/\Gamma(K^+ K^-)$ $\Gamma_{30}/\Gamma_{29}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.52±0.05 OUR FIT</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.70±0.21±0.12	<sup>1,2</sup> CHEN	07B	BELL $e^+ e^- \rightarrow e^+ e^- \chi_{c2}$

<sup>1</sup> Using  $\Gamma(K^+ K^-) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$  from NAKAZAWA 05.

<sup>2</sup> Not independent from other measurements.

### $\Gamma(K^+ K^- \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{32}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>0.32±0.08±0.01</b>	<sup>1</sup> ATHAR	07	CLEO $\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ATHAR 07 reports  $(0.31 \pm 0.07 \pm 0.04) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(K^+ K^- \eta)/\Gamma_{\text{total}}$ $\Gamma_{33}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.34</b>	90	<sup>1</sup> ATHAR	07	CLEO $\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ATHAR 07 reports  $< 0.33 \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

### $\Gamma(K^+ K^- \eta'(958))/\Gamma_{\text{total}}$ $\Gamma_{34}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.94±0.34</b>	107	<sup>1</sup> ABLIKIM	14J	BES3 $\psi(2S) \rightarrow \gamma K^+ K^- \eta'(958)$

<sup>1</sup> Derived using  $B(\psi(2S) \rightarrow \gamma\chi_{c2}) = (8.72 \pm 0.34)\%$ . Uncertainty includes both statistical and systematic contributions combined in quadrature.

### $\Gamma(\eta\eta')/\Gamma_{\text{total}}$ $\Gamma_{35}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.6</b>	90	3.3 ± 8.0	<sup>1</sup> ASNER	09	CLEO $\psi(2S) \rightarrow \gamma \eta \eta'$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<2.4	90	<sup>2</sup> ADAMS	07	CLEO $\psi(2S) \rightarrow \gamma\chi_{c2}$
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<sup>1</sup> ASNER 09 reports  $< 0.6 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \eta\eta')/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

<sup>2</sup> Superseded by ASNER 09. ADAMS 07 reports  $< 2.3 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \eta\eta')/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 0.0933 \pm 0.0014 \pm 0.0061$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

### $\Gamma(\eta'\eta')/\Gamma_{\text{total}}$ $\Gamma_{36}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0</b>	90	12 ± 7	<sup>1</sup> ASNER	09	CLEO $\psi(2S) \rightarrow \gamma \eta' \eta'$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<3.2	90	<sup>2</sup> ADAMS	07	CLEO $\psi(2S) \rightarrow \gamma\chi_{c2}$
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<sup>1</sup> ASNER 09 reports  $< 1.0 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \eta'\eta')/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

<sup>2</sup> Superseded by ASNER 09. ADAMS 07 reports  $< 3.1 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \eta'\eta')/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 0.0933 \pm 0.0014 \pm 0.0061$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

### $\Gamma(\pi^+ \pi^- K_S^0 K_S^0)/\Gamma_{\text{total}}$ $\Gamma_{37}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.3±0.6±0.1</b>	57 ± 11	<sup>1</sup> ABLIKIM	05o	BES2 $\psi(2S) \rightarrow \gamma\chi_{c2}$

<sup>1</sup> ABLIKIM 05o reports  $[\Gamma(\chi_{c2}(1P) \rightarrow \pi^+ \pi^- K_S^0 K_S^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  =  $(0.207 \pm 0.039 \pm 0.033) \times 10^{-3}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(K^+ K^- K_S^0 K_S^0)/\Gamma_{\text{total}}$ $\Gamma_{38}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;4</b>	90	2.3 ± 2.2	<sup>1</sup> ABLIKIM	05o	BES2 $e^+ e^- \rightarrow \chi_{c2} \gamma$

<sup>1</sup> ABLIKIM 05o reports  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- K_S^0 K_S^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  <  $3.5 \times 10^{-5}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

### $\Gamma(K^+ K^- K^+ K^-)/\Gamma_{\text{total}}$ $\Gamma_{39}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID
<b>1.73±0.21 OUR FIT</b>	

### $\Gamma(K^+ K^- \phi)/\Gamma_{\text{total}}$ $\Gamma_{40}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.48±0.31±0.05</b>	52	<sup>1</sup> ABLIKIM	06T	BES2 $\psi(2S) \rightarrow \gamma 2K^+ 2K^-$

<sup>1</sup> ABLIKIM 06T reports  $(1.67 \pm 0.26 \pm 0.24) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- \phi)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.1 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(K^0 \bar{K}^0 \pi^- \phi + \text{c.c.})/\Gamma_{\text{total}}$ $\Gamma_{41}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>4.83±0.32±0.66</b>	ABLIKIM	15M	BES3 $\psi(2S) \rightarrow \gamma\chi_{c2}$

### $\Gamma(K^+ K^- \pi^0 \phi)/\Gamma_{\text{total}}$ $\Gamma_{42}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>2.74±0.16±0.44</b>	ABLIKIM	15M	BES3 $\psi(2S) \rightarrow \gamma\chi_{c2}$

### $\Gamma(\phi \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{43}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.93±0.06±0.10</b>	408	<sup>1</sup> ABLIKIM	13B	BES3 $e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma\chi_{c2}$

<sup>1</sup> Using  $1.06 \times 10^8$   $\psi(2S)$  mesons and  $B(\psi(2S) \rightarrow \chi_{c2} \gamma) = (8.72 \pm 0.34)\%$ .

### $\Gamma(p\bar{p})/\Gamma_{\text{total}}$ $\Gamma_{44}/\Gamma$

VALUE (units 10 <sup>-4</sup> )	DOCUMENT ID
<b>0.75±0.04 OUR FIT</b>	

### $\Gamma(p\bar{p} \pi^0)/\Gamma_{\text{total}}$ $\Gamma_{45}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>0.49±0.04 OUR AVERAGE</b>			
0.49±0.04±0.02	<sup>1</sup> ONYISI	10	CLE3 $\psi(2S) \rightarrow \gamma p \bar{p} X$
0.45±0.09±0.02	<sup>2</sup> ATHAR	07	CLEO $\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ONYISI 10 reports  $(4.83 \pm 0.25 \pm 0.35 \pm 0.31) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p} \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ATHAR 07 reports  $(0.44 \pm 0.08 \pm 0.05) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p} \pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(p\bar{p} \eta)/\Gamma_{\text{total}}$ $\Gamma_{46}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>0.182±0.026 OUR AVERAGE</b>			
0.180±0.027±0.006	<sup>1</sup> ONYISI	10	CLE3 $\psi(2S) \rightarrow \gamma p \bar{p} X$
0.19 ± 0.07 ± 0.01	<sup>2</sup> ATHAR	07	CLEO $\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup> ONYISI 10 reports  $(1.76 \pm 0.23 \pm 0.14 \pm 0.11) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p} \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> ATHAR 07 reports  $(0.19 \pm 0.07 \pm 0.02) \times 10^{-3}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p} \eta)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(p\bar{p} \omega)/\Gamma_{\text{total}}$ $\Gamma_{47}/\Gamma$

VALUE (units 10 <sup>-3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>0.38±0.04±0.01</b>	<sup>1</sup> ONYISI	10	CLE3 $\psi(2S) \rightarrow \gamma p \bar{p} X$

<sup>1</sup> ONYISI 10 reports  $(3.68 \pm 0.35 \pm 0.26 \pm 0.24) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p} \omega)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(p\bar{p} \phi)/\Gamma_{\text{total}}$ $\Gamma_{48}/\Gamma$

VALUE (units 10 <sup>-5</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.9±0.9±0.1</b>	24 ± 7	<sup>1</sup> ABLIKIM	11F	BES3 $\psi(2S) \rightarrow \gamma p \bar{p} K^+ K^-$

<sup>1</sup> ABLIKIM 11F reports  $(3.04 \pm 0.85 \pm 0.43) \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p} \phi)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

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## Meson Particle Listings

 $\chi_{c2}(1P)$  $\Gamma(p\bar{p}\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{49}/\Gamma$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.32±0.34 OUR EVALUATION</b>	Treating systematic error as correlated.		
<b>1.3 ±0.4 OUR AVERAGE</b>	Error includes scale factor of 1.3.		
1.17±0.19±0.30	<sup>1</sup> BAI	99B BES	$\psi(2S) \rightarrow \gamma\chi_{c2}$
2.64±1.03±0.14	<sup>1</sup> TANENBAUM	78 MRK1	$\psi(2S) \rightarrow \gamma\chi_{c2}$

<sup>1</sup>Rescaled by us using  $B(\psi(2S) \rightarrow \gamma\chi_{c2}) = (8.3 \pm 0.4)\%$  and  $B(\psi(2S) \rightarrow J/\psi(1S)\pi^+\pi^-) = (32.6 \pm 0.5)\%$ . Multiplied by a factor of 2 to convert from  $K_S^0 K^+\pi^-$  to  $K^0 K^+\pi^-$  decay.

 $\Gamma(p\bar{p}\pi^0\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{50}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.082±0.024±0.003</b>	29.2	<sup>1</sup> HE	08B CLEO	$e^+e^- \rightarrow \gamma h^+ h^- h^0 h^0$

<sup>1</sup>HE 08B reports  $0.08 \pm 0.02 \pm 0.01 \pm 0.01\%$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p}\pi^0\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(p\bar{p}K^+K^- \text{ (non-resonant)})/\Gamma_{\text{total}}$   $\Gamma_{51}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.08±0.33±0.07</b>	131 ± 12	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\bar{p}K^+K^-$

<sup>1</sup>ABLIKIM 11F reports  $(2.08 \pm 0.19 \pm 0.30) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{p}K^+K^- \text{ (non-resonant)})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(p\bar{p}K_S^0K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{52}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7.9</b>	90	<sup>1</sup> ABLIKIM	06D BES2	$\psi(2S) \rightarrow \chi_{c2}\gamma$

<sup>1</sup>Using  $B(\psi(2S) \rightarrow \chi_{c2}\gamma) = (9.3 \pm 0.6)\%$ .

 $\Gamma(p\bar{\pi}\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{53}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.9±1.0 OUR AVERAGE</b>				
8.8±1.0±0.3	3309	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma p\bar{\pi}\pi^-$
10.6±3.6±0.4		<sup>2</sup> ABLIKIM	06I BES2	$\psi(2S) \rightarrow \gamma p\pi^- X$

<sup>1</sup>ABLIKIM 12J reports  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{\pi}\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  =  $(0.80 \pm 0.02 \pm 0.09) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup>ABLIKIM 06I reports  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{\pi}\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  =  $(0.97 \pm 0.20 \pm 0.26) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\bar{p}n\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{54}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.3±0.8±0.3</b>	3732	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma\bar{p}n\pi^+$

<sup>1</sup>ABLIKIM 12J reports  $[\Gamma(\chi_{c2}(1P) \rightarrow \bar{p}n\pi^+)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  =  $(0.85 \pm 0.02 \pm 0.07) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(p\bar{\pi}\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{55}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>22.7±1.8±0.8</b>	2128	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma p\bar{\pi}\pi^-\pi^0$

<sup>1</sup>ABLIKIM 12J reports  $[\Gamma(\chi_{c2}(1P) \rightarrow p\bar{\pi}\pi^-\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  =  $(2.07 \pm 0.06 \pm 0.15) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\bar{p}n\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{56}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>22.1±1.9±0.8</b>	2352	<sup>1</sup> ABLIKIM	12J BES3	$\psi(2S) \rightarrow \gamma\bar{p}n\pi^+\pi^0$

<sup>1</sup>ABLIKIM 12J reports  $[\Gamma(\chi_{c2}(1P) \rightarrow \bar{p}n\pi^+\pi^0)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  =  $(2.01 \pm 0.06 \pm 0.16) \times 10^{-4}$  which we divide by our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\Lambda\bar{\Lambda})/\Gamma_{\text{total}}$   $\Gamma_{57}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>1.92±0.16 OUR FIT</b>	

 $\Gamma(\Lambda\bar{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{58}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>131±16±5</b>		371	<sup>1</sup> ABLIKIM	12I BES3	$\psi(2S) \rightarrow \gamma\Lambda\bar{\Lambda}\pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<350 90 <sup>2</sup>ABLIKIM 06D BES2  $\psi(2S) \rightarrow \chi_{c2}\gamma$

<sup>1</sup>ABLIKIM 12I reports  $(137.0 \pm 7.6 \pm 15.7) \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Lambda\bar{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup>Using  $B(\psi(2S) \rightarrow \chi_{c2}\gamma) = (9.3 \pm 0.6)\%$ .

 $\Gamma(\Lambda\bar{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}$   $\Gamma_{59}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>69±16±2</b>		36	<sup>1</sup> ABLIKIM	12I BES3	$\psi(2S) \rightarrow \gamma\Lambda\bar{\Lambda}\pi^+\pi^-$

<sup>1</sup>ABLIKIM 12I reports  $(71.8 \pm 14.5 \pm 8.2) \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Lambda\bar{\Lambda}\pi^+\pi^- \text{ (non-resonant)})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\Sigma(1385)^+\bar{\Lambda}\pi^- + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{60}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;40</b>	90	<sup>1</sup> ABLIKIM	12I BES3	$\psi(2S) \rightarrow \gamma\Sigma(1385)^+\bar{\Lambda}\pi^-$

<sup>1</sup>ABLIKIM 12I reports  $< 42 \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma(1385)^+\bar{\Lambda}\pi^- + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

 $\Gamma(\Sigma(1385)^-\bar{\Lambda}\pi^+ + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{61}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;60</b>	90	<sup>1</sup> ABLIKIM	12I BES3	$\psi(2S) \rightarrow \gamma\Sigma(1385)^-\bar{\Lambda}\pi^+$

<sup>1</sup>ABLIKIM 12I reports  $< 61 \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma(1385)^-\bar{\Lambda}\pi^+ + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

 $\Gamma(K^+\bar{p}\Lambda + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{62}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.1±0.6 OUR AVERAGE</b>					
8.0±0.6±0.3		5k	<sup>1,2</sup> ABLIKIM	13D BES3	$\psi(2S) \rightarrow \gamma\Lambda\bar{p}K^+$
8.7±1.7±0.3			<sup>3</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma h^+ h^- h^0$

<sup>1</sup>ABLIKIM 13D reports  $(8.4 \pm 0.3 \pm 0.6) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+\bar{p}\Lambda + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup>Using  $B(\Lambda \rightarrow p\pi^-) = 63.9\%$ .

<sup>3</sup>ATHAR 07 reports  $(8.5 \pm 1.4 \pm 1.0) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+\bar{p}\Lambda + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(K^+\bar{p}\Lambda(1520) + \text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{63}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.9±0.7±0.1</b>		79 ± 13	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\bar{p}K^+K^-$

<sup>1</sup>ABLIKIM 11F reports  $(3.06 \pm 0.50 \pm 0.54) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow K^+\bar{p}\Lambda(1520) + \text{c.c.})/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\Lambda(1520)\bar{\Lambda}(1520))/\Gamma_{\text{total}}$   $\Gamma_{64}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.8±1.5±0.2</b>		29 ± 7	<sup>1</sup> ABLIKIM	11F BES3	$\psi(2S) \rightarrow \gamma p\bar{p}K^+K^-$

<sup>1</sup>ABLIKIM 11F reports  $(5.05 \pm 1.29 \pm 0.93) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Lambda(1520)\bar{\Lambda}(1520))/\Gamma_{\text{total}}] \times [B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $B(\psi(2S) \rightarrow \gamma\chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\Sigma^0\bar{\Sigma}^0)/\Gamma_{\text{total}}$   $\Gamma_{65}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.6</b>		90	<sup>1</sup> ABLIKIM	13H BES3	$\psi(2S) \rightarrow \gamma\Sigma^0\bar{\Sigma}^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.8 90 7.5 ± 3.4 <sup>2</sup>NAIK 08 CLEO  $\psi(2S) \rightarrow \gamma\Sigma^0\bar{\Sigma}^0$

# Meson Particle Listings

## $\chi_{c2}(1P)$

<sup>1</sup> ABLIKIM 13H reports  $< 0.65 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma^0 \bar{\Xi}^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .  
<sup>2</sup> NAIK 08 reports  $< 0.75 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma^0 \bar{\Xi}^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\Sigma^+ \bar{\Sigma}^-)/\Gamma_{\text{total}}$		$\Gamma_{66}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
<0.7	90	$4.0 \pm 3.5$	<sup>1</sup> NAIK 08 CLEO $\psi(2S) \rightarrow \gamma \Sigma^+ \bar{\Sigma}^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.8	90		<sup>2</sup> ABLIKIM 13H BES3 $\psi(2S) \rightarrow \gamma \Sigma^+ \bar{\Sigma}^-$

<sup>1</sup> NAIK 08 reports  $< 0.67 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma^+ \bar{\Sigma}^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .  
<sup>2</sup> ABLIKIM 13H reports  $< 0.88 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma^+ \bar{\Sigma}^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\Sigma(1385)^+ \bar{\Sigma}(1385)^-)/\Gamma_{\text{total}}$		$\Gamma_{67}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<16	90	<sup>1</sup> ABLIKIM 12i BES3	$\psi(2S) \rightarrow \gamma \Sigma(1385)^+ \bar{\Sigma}(1385)^-$

<sup>1</sup> ABLIKIM 12i reports  $< 17 \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma(1385)^+ \bar{\Sigma}(1385)^-)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\Sigma(1385)^- \bar{\Sigma}(1385)^+)/\Gamma_{\text{total}}$		$\Gamma_{68}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
<8	90	<sup>1</sup> ABLIKIM 12i BES3	$\psi(2S) \rightarrow \gamma \Sigma(1385)^- \bar{\Sigma}(1385)^+$

<sup>1</sup> ABLIKIM 12i reports  $< 8.5 \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Sigma(1385)^- \bar{\Sigma}(1385)^+)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.72 \pm 0.34) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(K^- \Lambda \bar{\Xi}^+ + \text{c.c.})/\Gamma_{\text{total}}$		$\Gamma_{69}/\Gamma$	
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b><math>1.84 \pm 0.33 \pm 0.06</math></b>	51	<sup>1</sup> ABLIKIM 15i BES3	$\psi(2S) \rightarrow \gamma K^- \Lambda \bar{\Xi}^+ + \text{c.c.}$

<sup>1</sup> ABLIKIM 15i reports  $[\Gamma(\chi_{c2}(1P) \rightarrow K^- \Lambda \bar{\Xi}^+ + \text{c.c.})/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  =  $(1.68 \pm 0.26 \pm 0.15) \times 10^{-5}$  which we divide by our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Xi^0 \bar{\Xi}^0)/\Gamma_{\text{total}}$		$\Gamma_{70}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
<1.1	90	$2.9 \pm 1.7$	<sup>1</sup> NAIK 08 CLEO $\psi(2S) \rightarrow \gamma \Xi^0 \bar{\Xi}^0$

<sup>1</sup> NAIK 08 reports  $< 1.06 \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Xi^0 \bar{\Xi}^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\Xi^- \bar{\Xi}^+)/\Gamma_{\text{total}}$		$\Gamma_{71}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
<b><math>1.48 \pm 0.33 \pm 0.05</math></b>		$29 \pm 5$	<sup>1</sup> NAIK 08 CLEO $\psi(2S) \rightarrow \gamma \Xi^- \bar{\Xi}^+$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
< 3.7 90 <sup>2</sup> ABLIKIM 06d BES2  $\psi(2S) \rightarrow \chi_{c2} \gamma$

<sup>1</sup> NAIK 08 reports  $(1.45 \pm 0.30 \pm 0.15) \times 10^{-4}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \Xi^- \bar{\Xi}^+)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.33 \pm 0.14 \pm 0.61) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (9.11 \pm 0.31) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>2</sup> Using  $\text{B}(\psi(2S) \rightarrow \chi_{c2} \gamma) = (9.3 \pm 0.6)\%$ .

$\Gamma(J/\psi(1S) \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$		$\Gamma_{72}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<0.015	90	BARATE 81 SPEC	$190 \text{ GeV } \pi^- \text{Be} \rightarrow 2\pi 2\mu$

$\Gamma(\pi^0 \eta_c)/\Gamma_{\text{total}}$		$\Gamma_{73}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<3.2 $\times 10^{-3}$	90	<sup>1</sup> ABLIKIM 15N BES3	$\psi(2S) e^+ e^- \rightarrow \gamma \pi^0 \eta_c$
<sup>1</sup> Using $\text{B}(\eta_c \rightarrow K_S^0 K^\pm \pi^\mp) \times \text{B}(K_S^0 \rightarrow \pi^+ \pi^-) \times \text{B}(\pi^0 \rightarrow \gamma) = (1.66 \pm 0.11) \times 10^{-2}$ .			

$\Gamma(\eta_c(1S) \pi^+ \pi^-)/\Gamma_{\text{total}}$		$\Gamma_{74}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt; 0.54 \times 10^{-2}</math></b>	90	<sup>1,2</sup> ABLIKIM 13B BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<1.2 $\times 10^{-2}$	90	<sup>1,3</sup> ABLIKIM 13B BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$
<sup>1</sup> Using $1.06 \times 10^8 \psi(2S)$ mesons and $\text{B}(\psi(2S) \rightarrow \chi_{c2} \gamma) = (8.72 \pm 0.34)\%$ .			
<sup>2</sup> From the $\eta_c \rightarrow K_S^0 K^\pm \pi^\mp$ decays.			
<sup>3</sup> From the $\eta_c \rightarrow K^+ K^- \pi^0$ decays.			

$\Gamma(\eta_c(1S) \pi^+ \pi^-)/\Gamma(\bar{K}^0 K^+ \pi^- + \text{c.c.})$		$\Gamma_{74}/\Gamma_{31}$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<16.4	90	<sup>1</sup> LEES	$12\text{AE BABR } e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \eta_c$
<sup>1</sup> We divided the reported limit by 2 to take into account the $K_L^0 K^+ \pi^-$ mode.			

## RADIATIVE DECAYS

$\Gamma(\gamma J/\psi(1S))/\Gamma_{\text{total}}$		$\Gamma_{75}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.192 \pm 0.007</math> OUR FIT</b>			• • • We do not use the following data for averages, fits, limits, etc. • • •
0.199 $\pm 0.005 \pm 0.012$	<sup>1</sup> ADAM 05A CLEO		$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$
<sup>1</sup> Uses $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2} \rightarrow \gamma \gamma J/\psi)$ from ADAM 05A and $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2})$ from ATHAR 04.			

$\Gamma(\gamma \rho^0)/\Gamma_{\text{total}}$		$\Gamma_{76}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
<20	90	$13 \pm 11$	<sup>1</sup> ABLIKIM 11E BES3 $\psi(2S) \rightarrow \gamma \gamma \rho^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<40	90	$17.2 \pm 6.8$	<sup>2</sup> BENNETT 08A CLEO $\psi(2S) \rightarrow \gamma \gamma \rho^0$

<sup>1</sup> ABLIKIM 11E reports  $< 20.8 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \gamma \rho^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .  
<sup>2</sup> BENNETT 08A reports  $< 50 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \gamma \rho^0)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.1 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\gamma \omega)/\Gamma_{\text{total}}$		$\Gamma_{77}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
<6	90	$1 \pm 6$	<sup>1</sup> ABLIKIM 11E BES3 $\psi(2S) \rightarrow \gamma \gamma \omega$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<6	90	$0.0 \pm 1.8$	<sup>2</sup> BENNETT 08A CLEO $\psi(2S) \rightarrow \gamma \gamma \omega$

<sup>1</sup> ABLIKIM 11E reports  $< 6.1 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \gamma \omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .  
<sup>2</sup> BENNETT 08A reports  $< 7.0 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \gamma \omega)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.1 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\gamma \phi)/\Gamma_{\text{total}}$		$\Gamma_{78}/\Gamma$	
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
< 8	90	$5 \pm 5$	<sup>1</sup> ABLIKIM 11E BES3 $\psi(2S) \rightarrow \gamma \gamma \phi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<12	90	$1.3 \pm 2.5$	<sup>2</sup> BENNETT 08A CLEO $\psi(2S) \rightarrow \gamma \gamma \phi$

<sup>1</sup> ABLIKIM 11E reports  $< 8.1 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \gamma \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .  
<sup>2</sup> BENNETT 08A reports  $< 13 \times 10^{-6}$  from a measurement of  $[\Gamma(\chi_{c2}(1P) \rightarrow \gamma \phi)/\Gamma_{\text{total}}] \times [\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))]$  assuming  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.1 \pm 0.4) \times 10^{-2}$ , which we rescale to our best value  $\text{B}(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = 9.11 \times 10^{-2}$ .

$\Gamma(\gamma \gamma)/\Gamma_{\text{total}}$		$\Gamma_{79}/\Gamma$	
VALUE (units $10^{-4}$ )	DOCUMENT ID		
<b><math>2.74 \pm 0.14</math> OUR FIT</b>			

$\Gamma(\gamma \gamma)/\Gamma(\gamma J/\psi(1S))$		$\Gamma_{79}/\Gamma_{75}$	
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.43 \pm 0.08</math> OUR FIT</b>			
<b><math>0.99 \pm 0.18</math></b>	<sup>1</sup> AMBROGIANI 00B E835		$\bar{p} p \rightarrow \chi_{c2} \rightarrow \gamma \gamma, \gamma J/\psi$
<sup>1</sup> Calculated by us using $\text{B}(J/\psi(1S) \rightarrow e^+ e^-) = 0.0593 \pm 0.0010$ .			

$\Gamma(\gamma \gamma)/\Gamma_{\text{total}} \times \Gamma(\bar{p} \bar{p})/\Gamma_{\text{total}}$		$\Gamma_{79}/\Gamma \times \Gamma_{44}/\Gamma$	
VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.06 \pm 0.16</math> OUR FIT</b>			
<b><math>1.7 \pm 0.4</math> OUR AVERAGE</b>			
1.60 $\pm 0.42$	ARMSTRONG 93 E760		$\bar{p} p \rightarrow \gamma \gamma X$
9.9 $\pm 4.5$	BAGLIN 87B SPEC		$\bar{p} p \rightarrow \gamma \gamma X$

## $\chi_{c2}(1P)$ CROSS-PARTICLE BRANCHING RATIOS

$\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P))/\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$		$\Gamma_{14}/\Gamma \times \Gamma_{134}^{\psi(2S)}/\Gamma_{11}^{\psi(2S)}$	
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.34 \pm 0.26</math> OUR FIT</b>			
<b><math>2.5 \pm 0.9</math> OUR AVERAGE</b>			Error includes scale factor of 2.3.
1.90 $\pm 0.14 \pm 0.44$	BAI 99B BES		$\psi(2S) \rightarrow \gamma \chi_{c2}$
3.8 $\pm 0.67$	<sup>1</sup> TANENBAUM 78 MRK1		$\psi(2S) \rightarrow \gamma \chi_{c2}$
<sup>1</sup> The reported value is derived using $\text{B}(\psi(2S) \rightarrow \pi^+ \pi^- J/\psi) \times \text{B}(J/\psi \rightarrow \ell^+ \ell^-) = (4.6 \pm 0.7)\%$ . Calculated by us using $\text{B}(J/\psi \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .			

$$\Gamma(\chi_{c2}(1P) \rightarrow K^*(892)^0 \bar{K}^*(892)^0) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{18} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.2 ± 0.4 OUR FIT</b>			
<b>3.11 ± 0.36 ± 0.48</b>	ABLIKIM	04H BES2	$\psi(2S) \rightarrow \gamma \chi_{c2}$

$$\Gamma(\chi_{c2}(1P) \rightarrow p \bar{p}) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{44} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.98 ± 0.10 OUR FIT</b>			
<b>1.4 ± 1.1</b>	<sup>1</sup> BAI	98i BES	$\psi(2S) \rightarrow \gamma \chi_{c2} \rightarrow \gamma \bar{p} p$

<sup>1</sup> Calculated by us. The value for  $B(\chi_{c2} \rightarrow p \bar{p})$  reported in BAI 98i is derived using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (7.8 \pm 0.8)\%$  and  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (32.4 \pm 2.6)\%$  [BAI 98d].

$$\Gamma(\chi_{c2}(1P) \rightarrow p \bar{p}) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{44} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.85 ± 0.33 OUR FIT</b>				
<b>7.1 ± 0.5 OUR AVERAGE</b>	Error includes scale factor of 1.2.			
7.3 ± 0.4 ± 0.3	405	ABLIKIM	13v BES3	$\psi(2S) \rightarrow \gamma p \bar{p}$
7.2 ± 0.7 ± 0.4	121 ± 12	<sup>1</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma p \bar{p}$
4.4 ± 1.6 ± 0.6	14.3 ± 5.2 ± 4.7	BAI	04f BES	$\psi(2S) \rightarrow \gamma \chi_{c2}(1P) \rightarrow \gamma \bar{p} p$

<sup>1</sup> Calculated by us. NAIK 08 reports  $B(\chi_{c2} \rightarrow p \bar{p}) = (7.7 \pm 0.8 \pm 0.4 \pm 0.5) \times 10^{-5}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow \Lambda \bar{\Lambda}) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{57} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.5 ± 1.3 OUR FIT</b>				
<b>17.4 ± 1.4 OUR AVERAGE</b>				
18.2 ± 1.4 ± 0.9	207	<sup>1</sup> ABLIKIM	13h BES3	$\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda}$
15.9 ± 2.1 ± 1.0	71 ± 9	<sup>2</sup> NAIK	08 CLEO	$\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda}$

<sup>1</sup> Calculated by us. ABLIKIM 13h reports  $B(\chi_{c2} \rightarrow \Lambda \bar{\Lambda}) = (20.8 \pm 1.6 \pm 2.3) \times 10^{-5}$  from a measurement of  $B(\chi_{c2} \rightarrow \Lambda \bar{\Lambda}) \times B(\psi(2S) \rightarrow \gamma \chi_{c2})$  assuming  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (8.74 \pm 0.35)\%$ .

<sup>2</sup> Calculated by us. NAIK 08 reports  $B(\chi_{c2} \rightarrow \Lambda \bar{\Lambda}) = (17.0 \pm 2.2 \pm 1.1 \pm 1.1) \times 10^{-5}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow \Lambda \bar{\Lambda}) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{57} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.1 ± 0.4 OUR FIT</b>				
<b>7.1 ± 3.1 ± 1.3</b>	8.3 ± 3.7 ± 3.4	<sup>1</sup> BAI	03e BES	$\psi(2S) \rightarrow \gamma \Lambda \bar{\Lambda}$

<sup>1</sup> BAI 03e reports  $[B(\chi_{c2} \rightarrow \Lambda \bar{\Lambda}) B(\psi(2S) \rightarrow \gamma \chi_{c2}) / B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-)] \times [B^2(\Lambda \rightarrow \pi^- p) / B(J/\psi \rightarrow p \bar{p})] = (1.33^{+0.59}_{-0.55} \pm 0.25)\%$ . We calculate from this measurement the presented value using  $B(\Lambda \rightarrow \pi^- p) = (63.9 \pm 0.5)\%$  and  $B(J/\psi \rightarrow p \bar{p}) = (2.17 \pm 0.07) \times 10^{-3}$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow \pi \pi) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{24} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.12 ± 0.08 OUR FIT</b>				
<b>2.17 ± 0.09 OUR AVERAGE</b>				
2.19 ± 0.05 ± 0.15	4.5k	<sup>1</sup> ABLIKIM	10A BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$
2.23 ± 0.06 ± 0.10	2.5k	<sup>2</sup> ASNER	09 CLEO	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-$
1.90 ± 0.08 ± 0.20	0.8k	<sup>3</sup> ASNER	09 CLEO	$\psi(2S) \rightarrow \gamma \pi^0 \pi^0$

<sup>1</sup> Calculated by us. ABLIKIM 10A reports  $B(\chi_{c2} \rightarrow \pi^0 \pi^0) = (0.88 \pm 0.02 \pm 0.06 \pm 0.04) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (8.3 \pm 0.4)\%$ . We have multiplied the  $\pi^0 \pi^0$  measurement by 3 to obtain  $\pi \pi$ .

<sup>2</sup> Calculated by us. ASNER 09 reports  $B(\chi_{c2} \rightarrow \pi^+ \pi^-) = (1.59 \pm 0.04 \pm 0.07 \pm 0.10) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ . We have multiplied the  $\pi^+ \pi^-$  measurement by 3/2 to obtain  $\pi \pi$ .

<sup>3</sup> Calculated by us. ASNER 09 reports  $B(\chi_{c2} \rightarrow \pi^0 \pi^0) = (0.68 \pm 0.03 \pm 0.07 \pm 0.04) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ . We have multiplied the  $\pi^0 \pi^0$  measurement by 3 to obtain  $\pi \pi$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow \pi \pi) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{24} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.615 ± 0.023 OUR FIT</b>				
<b>0.54 ± 0.06 OUR AVERAGE</b>				
0.66 ± 0.18 ± 0.37	21 ± 6	<sup>1</sup> BAI	03c BES	$\psi(2S) \rightarrow \gamma \pi^0 \pi^0$
0.54 ± 0.05 ± 0.04	185 ± 16	<sup>2</sup> BAI	98i BES	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-$

<sup>1</sup> We have multiplied  $\pi^0 \pi^0$  measurement by 3 to obtain  $\pi \pi$ .

<sup>2</sup> Calculated by us. The value for  $B(\chi_{c2} \rightarrow \pi^+ \pi^-)$  reported by BAI 98i is derived using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (7.8 \pm 0.8)\%$  and  $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$  [BAI 98d]. We have multiplied  $\pi^+ \pi^-$  measurement by 3/2 to obtain  $\pi \pi$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow \eta \eta) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{28} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-4}$ )	CL% EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.52 ± 0.04 OUR FIT</b>				
<b>0.52 ± 0.04 OUR AVERAGE</b>				
0.54 ± 0.03 ± 0.04	386	<sup>1</sup> ABLIKIM	10A BES3	$e^+ e^- \rightarrow \psi(2S) \rightarrow \gamma \chi_{c2}$
0.47 ± 0.05 ± 0.05	156	ASNER	09 CLEO	$\psi(2S) \rightarrow \gamma \eta \eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.44	90	<sup>2</sup> ADAMS	07 CLEO	$\psi(2S) \rightarrow \gamma \chi_{c2}$
< 3	90	BAI	03c BES	$\psi(2S) \rightarrow \gamma \eta \eta \rightarrow 5 \gamma$
0.62 ± 0.31 ± 0.19		LEE	85 CBAL	$\psi(2S) \rightarrow \text{photons}$

<sup>1</sup> Calculated by us. ABLIKIM 10A reports  $B(\chi_{c2} \rightarrow \eta \eta) = (0.65 \pm 0.04 \pm 0.05 \pm 0.03) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (8.3 \pm 0.4)\%$ .

<sup>2</sup> Superseded by ASNER 09.

$$\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^-) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{29} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.6 ± 0.6 OUR FIT</b>				
<b>10.5 ± 0.3 ± 0.6</b>	1.6k	<sup>1</sup> ASNER	09 CLEO	$\psi(2S) \rightarrow \gamma K^+ K^-$

<sup>1</sup> Calculated by us. ASNER 09 reports  $B(\chi_{c2} \rightarrow K^+ K^-) = (1.13 \pm 0.03 \pm 0.06 \pm 0.07) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^-) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{29} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.277 ± 0.017 OUR FIT</b>				
<b>0.190 ± 0.034 ± 0.019</b>	115 ± 13	<sup>1</sup> BAI	98i BES	$\psi(2S) \rightarrow \gamma K^+ K^-$

<sup>1</sup> Calculated by us. The value for  $B(\chi_{c2} \rightarrow K^+ K^-)$  reported by BAI 98i is derived using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (7.8 \pm 0.8)\%$  and  $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$  [BAI 98d].

$$\Gamma(\chi_{c2}(1P) \rightarrow K_S^0 K_S^0) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{30} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.0 ± 0.4 OUR FIT</b>				
<b>5.0 ± 0.4 OUR AVERAGE</b>				
4.9 ± 0.3 ± 0.3	373 ± 20	<sup>1</sup> ASNER	09 CLEO	$\psi(2S) \rightarrow \gamma K_S^0 K_S^0$
5.72 ± 0.76 ± 0.63	65	ABLIKIM	05o BES2	$\psi(2S) \rightarrow \gamma K_S^0 K_S^0$

<sup>1</sup> Calculated by us. ASNER 09 reports  $B(\chi_{c2} \rightarrow K_S^0 K_S^0) = (0.53 \pm 0.03 \pm 0.03 \pm 0.03) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow K_S^0 K_S^0) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_{30} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>14.5 ± 1.1 OUR FIT</b>			
<b>14.7 ± 4.1 ± 3.3</b>	<sup>1</sup> BAI	99b BES	$\psi(2S) \rightarrow \gamma K_S^0 K_S^0$

<sup>1</sup> Calculated by us. The value of  $B(\chi_{c2} \rightarrow K_S^0 K_S^0)$  reported by BAI 99b was derived using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (7.8 \pm 0.8)\%$  and  $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$  [BAI 98d].

$$\Gamma(\chi_{c2}(1P) \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.}) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \times \Gamma_{31} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma \psi(2S)$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.22 ± 0.17 OUR FIT</b>				
<b>1.15 ± 0.18 OUR AVERAGE</b>				
1.21 ± 0.19 ± 0.09	37	<sup>1</sup> ATHAR	07 CLEO	$\psi(2S) \rightarrow \gamma K_S^0 K^{\pm} \pi^{\mp}$
0.97 ± 0.32 ± 0.13	28	<sup>2</sup> ABLIKIM	06r BES2	$\psi(2S) \rightarrow \gamma K_S^0 K^{\pm} \pi^{\mp}$

<sup>1</sup> Calculated by us. ATHAR 07 reports  $B(\chi_{c2} \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.}) = (1.3 \pm 0.2 \pm 0.1 \pm 0.1) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (9.33 \pm 0.14 \pm 0.61)\%$ .

<sup>2</sup> Calculated by us. ABLIKIM 06r reports  $B(\chi_{c2} \rightarrow K_S^0 K^{\pm} \pi^{\mp}) = (0.6 \pm 0.2 \pm 0.1) \times 10^{-3}$  using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (8.1 \pm 0.6)\%$ . We have multiplied by 2 to obtain  $\bar{K}^0 K^+ \pi^- + \text{c.c.}$  from  $K_S^0 K^{\pm} \pi^{\mp}$ .

$$\Gamma(\chi_{c2}(1P) \rightarrow 2(\pi^+ \pi^-)) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma_1 / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.83 ± 0.27 OUR FIT</b>			
<b>3.1 ± 1.0 OUR AVERAGE</b>	Error includes scale factor of 2.5.		
2.3 ± 0.1 ± 0.5	<sup>1</sup> BAI	99b BES	$\psi(2S) \rightarrow \gamma \chi_{c2}$
4.3 ± 0.6	<sup>2</sup> TANENBAUM	78 MRK1	$\psi(2S) \rightarrow \gamma \chi_{c2}$

<sup>1</sup> Calculated by us. The value for  $B(\chi_{c2} \rightarrow 2\pi^+ 2\pi^-)$  reported in BAI 99b is derived using  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) = (7.8 \pm 0.8)\%$  and  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (32.4 \pm 2.6)\%$  [BAI 98d].

<sup>2</sup> The value for  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) \times B(\chi_{c2} \rightarrow 2\pi^+ 2\pi^-)$  reported in TANENBAUM 78 is derived using  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times B(J/\psi(1S) \ell^+ \ell^-) = (4.6 \pm 0.7)\%$ . Calculated by us using  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .

# Meson Particle Listings

## $\chi_{c2}(1P)$

$$\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- K^+ K^-) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}}$$

$$\Gamma_{39} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.57±0.19 OUR FIT</b>				
<b>1.76±0.16±0.24</b>	160	<sup>1</sup> ABLIKIM	06T BES2	$\psi(2S) \rightarrow \gamma 2K^+ 2K^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c2} \rightarrow 2K^+ 2K^-)$ reported by ABLIKIM 06T was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.1 \pm 0.4)\%$ .				

$$\Gamma(\chi_{c2}(1P) \rightarrow K^+ K^- K^+ K^-) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$$

$$\Gamma_{39} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>4.6±0.5 OUR FIT</b>			
<b>3.6±0.6±0.6</b>	<sup>1</sup> BAI	99B BES	$\psi(2S) \rightarrow \gamma 2K^+ 2K^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c2} \rightarrow 2K^+ 2K^-)$ reported by BAI 99B was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (7.8 \pm 0.8)\%$ and $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$ [BAI 98d].			

$$\Gamma(\chi_{c2}(1P) \rightarrow \phi \phi) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}}$$

$$\Gamma_{20} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.02±0.08 OUR FIT</b>				
<b>0.98±0.13 OUR AVERAGE</b>				Error includes scale factor of 1.3.
0.94±0.03±0.10	849	<sup>1</sup> ABLIKIM	11K BES3	$\psi(2S) \rightarrow \gamma$ hadrons
1.38±0.24±0.23	41	<sup>2</sup> ABLIKIM	06T BES2	$\psi(2S) \rightarrow \gamma 2K^+ 2K^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c2} \rightarrow \phi \phi)$ reported by ABLIKIM 11K was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.74 \pm 0.35)\%$ .				
<sup>2</sup> Calculated by us. The value of $B(\chi_{c2} \rightarrow \phi \phi)$ reported by ABLIKIM 06T was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (8.1 \pm 0.4)\%$ .				

$$\Gamma(\chi_{c2}(1P) \rightarrow \phi \phi) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$$

$$\Gamma_{20} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>2.95±0.24 OUR FIT</b>			
<b>4.8 ± 1.3 ± 1.3</b>	<sup>1</sup> BAI	99B BES	$\psi(2S) \rightarrow \gamma 2K^+ 2K^-$
<sup>1</sup> Calculated by us. The value of $B(\chi_{c2} \rightarrow \phi \phi)$ reported by BAI 99B was derived using $B(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) = (7.8 \pm 0.8)\%$ and $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = (32.4 \pm 2.6)\%$ [BAI 98d].			

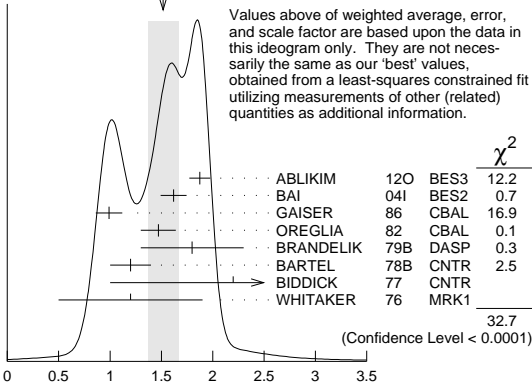
$$\Gamma(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}}$$

$$\Gamma_{75} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma^{\psi(2S)}$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.75 ± 0.04 OUR FIT</b>				
<b>1.52 ± 0.15 OUR AVERAGE</b>				Error includes scale factor of 2.6. See the ideogram below.
1.874±0.007±0.102	76k	ABLIKIM	120 BES3	$\psi(2S) \rightarrow \gamma \chi_{c2}$
1.62 ± 0.04 ± 0.12	5.8k	BAI	04I BES2	$\psi(2S) \rightarrow J/\psi \gamma \gamma$
0.99 ± 0.10 ± 0.08		GAISER	86 CBAL	$\psi(2S) \rightarrow \gamma X$
1.47 ± 0.17		<sup>1</sup> OREGLIA	82 CBAL	$\psi(2S) \rightarrow \gamma \chi_{c2}$
1.8 ± 0.5		<sup>2</sup> BRANDELIK	79B DASP	$\psi(2S) \rightarrow \gamma \chi_{c2}$
1.2 ± 0.2		<sup>2</sup> BARTEL	78B CNTR	$\psi(2S) \rightarrow \gamma \chi_{c2}$
2.2 ± 1.2		<sup>3</sup> BIDDICK	77 CNTR	$e^+ e^- \rightarrow \gamma X$
1.2 ± 0.7		<sup>1</sup> WHITAKER	76 MRK1	$e^+ e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.95 ± 0.02 ± 0.07	12.4k	<sup>4</sup> MENDEZ	08 CLEO	$\psi(2S) \rightarrow \gamma \chi_{c2}$
1.85 ± 0.04 ± 0.07	1.9k	<sup>5</sup> ADAM	05A CLEO	Repl. by MENDEZ 08

- <sup>1</sup> Recalculated by us using  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .  
<sup>2</sup> Recalculated by us using  $B(J/\psi(1S) \rightarrow \mu^+ \mu^-) = 0.0588 \pm 0.0010$ .  
<sup>3</sup> Assumes isotropic gamma distribution.  
<sup>4</sup> Not independent from other measurements of MENDEZ 08.  
<sup>5</sup> Not independent from other values reported by ADAM 05A.

WEIGHTED AVERAGE  
1.52±0.15 (Error scaled by 2.6)



$$\Gamma(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}} \text{ (units } 10^{-2})$$

$$\Gamma(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \text{ anything})$$

$$\Gamma_{75} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_9^{\psi(2S)}$$

$$\Gamma_{75} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_9^{\psi(2S)} = \Gamma_{75} / \Gamma \times \Gamma_{134}^{\psi(2S)} / (\Gamma_{11}^{\psi(2S)} + \Gamma_{12}^{\psi(2S)} + \Gamma_{13}^{\psi(2S)} + 0.339 \Gamma_{133}^{\psi(2S)} + 0.192 \Gamma_{134}^{\psi(2S)})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.87±0.07 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.12±0.03±0.09	12.4k	<sup>1</sup> MENDEZ	08 CLEO	$\psi(2S) \rightarrow \gamma \chi_{c2}$
3.11±0.07±0.07	1.9k	ADAM	05A CLEO	Repl. by MENDEZ 08
<sup>1</sup> Not independent from other measurements of MENDEZ 08.				

$$\Gamma(\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)$$

$$\Gamma_{75} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma_{11}^{\psi(2S)}$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.08±0.12 OUR FIT</b>				
<b>5.53±0.17 OUR AVERAGE</b>				
5.56±0.05±0.16	12.4k	MENDEZ	08 CLEO	$\psi(2S) \rightarrow \gamma \chi_{c2}$
6.0 ± 2.8	1.3k	<sup>1</sup> ABLIKIM	04B BES	$\psi(2S) \rightarrow J/\psi X$
3.9 ± 1.2		<sup>2</sup> HIMEL	80 MRK2	$\psi(2S) \rightarrow \gamma \chi_{c2}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5.52±0.13±0.13	1.9k	<sup>3</sup> ADAM	05A CLEO	Repl. by MENDEZ 08

- <sup>1</sup> From a fit to the  $J/\psi$  recoil mass spectra.  
<sup>2</sup> The value for  $B(\psi(2S) \rightarrow \gamma \chi_{c2}) \times B(\chi_{c2} \rightarrow \gamma J/\psi(1S))$  reported in HIMEL 80 is derived using  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (33 \pm 3)\%$  and  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.138 \pm 0.018$ . Calculated by us using  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = (0.1181 \pm 0.0020)$ .  
<sup>3</sup> Not independent from other values reported by ADAM 05A.

$$\Gamma(\chi_{c2}(1P) \rightarrow \gamma \gamma) / \Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma \chi_{c2}(1P)) / \Gamma_{\text{total}}$$

$$\Gamma_{79} / \Gamma \times \Gamma_{134}^{\psi(2S)} / \Gamma^{\psi(2S)}$$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.50±0.13 OUR FIT</b>				
<b>2.78±0.18 OUR AVERAGE</b>				
2.81±0.17±0.15	1.1k	<sup>1</sup> ABLIKIM	12A BES3	$\psi(2S) \rightarrow \gamma \chi_{c2} \rightarrow 3\gamma$
2.68±0.28±0.15	0.3k	ECKLUND	08A CLEO	$\psi(2S) \rightarrow \gamma \chi_{c2} \rightarrow 3\gamma$
7.0 ± 2.1 ± 2.0		LEE	85 CBAL	$\psi(2S) \rightarrow \gamma \chi_{c2}$
<sup>1</sup> ABLIKIM 12A measures the ratio of two-photon partial widths for the helicity $\lambda = 0$ and helicity $\lambda = 2$ components to be $f_{0/2} = \Gamma_{\gamma\gamma}^{\lambda=0} / \Gamma_{\gamma\gamma}^{\lambda=2} = 0.00 \pm 0.02 \pm 0.02$ .				

$$\Gamma(\chi_{c2}(1P) \rightarrow \gamma \gamma) / \Gamma(\chi_{c0}(1P) \rightarrow \gamma \gamma)$$

$$\Gamma_{79} / \Gamma_{89}^{\chi_{c0}(1P)}$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.273±0.035 OUR AVERAGE</b>				
0.271±0.029±0.030	1.9k	<sup>1</sup> ABLIKIM	12A BES3	$\psi(2S) \rightarrow \gamma \chi_{cJ} \rightarrow 3\gamma$
0.278±0.050±0.036	0.5k	<sup>1</sup> ECKLUND	08A CLEO	$\psi(2S) \rightarrow \gamma \chi_{cJ} \rightarrow 3\gamma$

- <sup>1</sup> Not independent from the values of  $\Gamma(\chi_{c0}, \chi_{c2})$  and  $B(\psi(2S) \rightarrow \chi_{c0}, \chi_{c2})$ .

## MULTIPOLE AMPLITUDES IN $\chi_{c2}(1P) \rightarrow \gamma J/\psi(1S)$ RADIATIVE DECAY

$a_2 = M_2 / \sqrt{E_1^2 + M_2^2 + E_3^2}$  Magnetic quadrupole fractional transition amplitude

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-10.0± 1.5 OUR AVERAGE</b>				
-9.3 ± 1.6±0.3	19.8k	<sup>1</sup> ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
-9.3 ± 3.9±0.6	5.9k	<sup>2</sup> AMBROGIANI	02 E835	$p\bar{p} \rightarrow \chi_{c2} \rightarrow J/\psi \gamma$
-14 ± 6	1.9k	<sup>2</sup> ARMSTRONG	93E E760	$p\bar{p} \rightarrow \chi_{c2} \rightarrow J/\psi \gamma$
-33.3 ± 11.6±29.2	441	<sup>2</sup> OREGLIA	82 CBAL	$\psi(2S) \rightarrow \chi_{c1} \gamma \rightarrow J/\psi \gamma \gamma$

- • • We do not use the following data for averages, fits, limits, etc. • • •
- 7.9 ± 1.9±0.3 19.8k <sup>3</sup> ARTUSO 09 CLEO  $\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
- <sup>1</sup> From a fit with floating  $M_2$  amplitudes  $a_2$  and  $b_2$ , and fixed  $E_3$  amplitudes  $a_3=b_3=0$ .  
<sup>2</sup> Assuming  $a_3=0$ .  
<sup>3</sup> From a fit with floating  $M_2$  and  $E_3$  amplitudes  $a_2, b_2$ , and  $a_3$ , and  $b_3$ .

$a_3 = E_3 / \sqrt{E_1^2 + M_2^2 + E_3^2}$  Electric octupole fractional transition amplitude

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.6±1.3 OUR AVERAGE</b>				
1.7±1.4±0.3	19.8k	<sup>1</sup> ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
2.0±5.5±4.4	5908	AMBROGIANI	02 E835	$p\bar{p} \rightarrow \chi_{c2} \rightarrow J/\psi \gamma$
0 ± 6±5	1904	ARMSTRONG	93E E760	$p\bar{p} \rightarrow \chi_{c2} \rightarrow J/\psi \gamma$

- <sup>1</sup> From a fit with floating  $M_2$  and  $E_3$  amplitudes  $a_2, b_2$ , and  $a_3$ , and  $b_3$ .



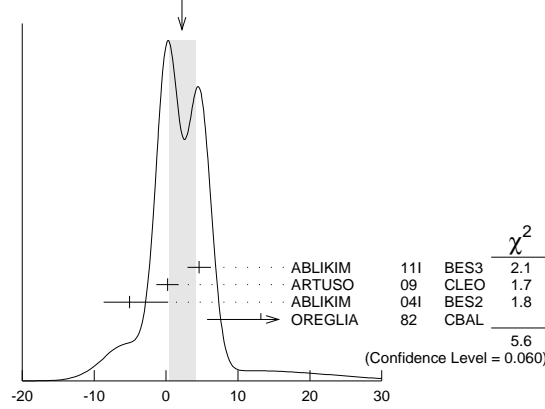
See key on page 601

## Meson Particle Listings

 $\chi_{c2}(1P), \eta_c(2S)$ MULTIPOLE AMPLITUDES IN  $\psi(2S) \rightarrow \gamma \chi_{c2}(1P)$  RADIATIVE DECAY $b_2 = M_2/\sqrt{E_1^2 + M_2^2 + E_3^2}$  Magnetic quadrupole fractional transition amplitude

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.2 \pm 1.8</math> OUR AVERAGE</b>		Error includes scale factor of 1.7. See the ideogram below.		
$4.6 \pm 1.0 \pm 1.3$	13.8k	<sup>1</sup> ABLIKIM	11i BES3	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-, \gamma K^+ K^-$
$0.2 \pm 1.5 \pm 0.4$	19.8k	<sup>2</sup> ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
$-5.1 \pm 5.4$ $-3.6$	721	<sup>1</sup> ABLIKIM	04i BES2	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-, \gamma K^+ K^-$
$13.2 \pm 9.8$ $7.5$	441	<sup>3</sup> OREGLIA	82 CBAL	$\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<sup>1</sup>  $1.0 \pm 1.3 \pm 0.3$  19.8k <sup>3</sup> ARTUSO 09 CLEO  $\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$ <sup>1</sup> From a fit with floating  $M_2$  and  $E_3$  amplitudes  $b_2$  and  $b_3$ .<sup>2</sup> From a fit with floating  $M_2$  and  $E_3$  amplitudes  $a_2, b_2$ , and  $a_3$ , and  $b_3$ .<sup>3</sup> From a fit with floating  $M_2$  amplitudes  $a_2$  and  $b_2$ , and fixed  $E_3$  amplitudes  $a_3=b_3=0$ .WEIGHTED AVERAGE  
 $2.2 \pm 1.8$  (Error scaled by 1.7) $b_2 = M_2/\sqrt{E_1^2 + M_2^2 + E_3^2}$  Magnetic quadrupole fractional transition amplitude (units  $10^{-2}$ ) $b_3 = E_3/\sqrt{E_1^2 + M_2^2 + E_3^2}$  Electric octupole fractional transition amplitude

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.3 \pm 1.0</math> OUR AVERAGE</b>				
$1.5 \pm 0.8 \pm 1.8$	13.8k	<sup>1</sup> ABLIKIM	11i BES3	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-, \gamma K^+ K^-$
$-0.8 \pm 1.2 \pm 0.2$	19.8k	ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
$-2.7 \pm 4.3$ $-2.9$	721	<sup>1</sup> ABLIKIM	04i BES2	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-, \gamma K^+ K^-$

<sup>1</sup> From a fit with floating  $M_2$  and  $E_3$  amplitudes  $b_2$  and  $b_3$ .

## MULTIPOLE AMPLITUDE RATIOS IN RADIATIVE DECAYS

 $\psi(2S) \rightarrow \gamma \chi_{c2}(1P)$  and  $\chi_{c2} \rightarrow \gamma J/\psi(1S)$  $b_2/a_2$  Magnetic quadrupole transition amplitude ratio

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-11 \pm 14</math> <math>-15</math></b>	19.8k	<sup>1</sup> ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$

<sup>1</sup> Statistical and systematic errors combined. From a fit with floating  $M_2$  amplitudes  $a_2$  and  $b_2$ , and fixed  $E_3$  amplitudes  $a_3=b_3=0$ . Not independent of values for  $a_2(\chi_{c2}(1P))$  and  $b_2(\chi_{c2}(1P))$  from ARTUSO 09. $\chi_{c2}(1P)$  REFERENCES

ABLIKIM	151	PR D91 092006	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	15M	PR D91 112008	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	15N	PR D91 112018	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	14J	PR D89 074030	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13B	PR D87 012002	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13D	PR D87 012007	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13H	PR D87 032007	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13V	PR D88 112001	M. Ablikim et al.	(BES III Collab.)
UEHARA	13	PTEP 2013 123C01	S. Uehara et al.	(BELLE Collab.)
ABLIKIM	12A	PR D85 112008	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	121	PR D86 052004	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12J	PR D86 052011	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12O	PRL 109 172002	M. Ablikim et al.	(BES III Collab.)
LEES	12AE	PR D86 092005	J.P. Lees et al.	(BABAR Collab.)
LIU	12B	PRL 108 232001	Z.Q. Liu et al.	(BELLE Collab.)
ABLIKIM	11A	PR D83 012006	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	11E	PR D83 112005	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	11F	PR D83 112009	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	111	PR D84 092006	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	11K	PRL 107 092001	M. Ablikim et al.	(BES III Collab.)
DEL-AMO-SA...	11M	PR D84 012004	P. del Amo Sanchez et al.	(BABAR Collab.)
ABLIKIM	10A	PR D81 052005	M. Ablikim et al.	(BES III Collab.)
ONYISI	10	PR D82 011103	P.U.E. Onyisi et al.	(CLEO Collab.)
UEHARA	10A	PR D82 114031	S. Uehara et al.	(BELLE Collab.)
ARTUSO	09	PR D80 112003	M. Artuso et al.	(CLEO Collab.)
ASNER	09	PR D79 072007	D.M. Asner et al.	(CLEO Collab.)

UEHARA	09	PR D79 052009	S. Uehara et al.	(BELLE Collab.)
BENNETT	08A	PRL 101 151801	J.V. Bennett et al.	(CLEO Collab.)
ECKLUND	08A	PR D78 091501	K.M. Ecklund et al.	(CLEO Collab.)
HE	08B	PR D78 092004	Q. He et al.	(CLEO Collab.)
MENDEZ	08	PR D78 011102	H. Mendez et al.	(CLEO Collab.)
NAIK	08	PR D78 031101	P. Naik et al.	(CLEO Collab.)
UEHARA	08	EPJ C53 1	S. Uehara et al.	(BELLE Collab.)
ADAMS	07	PR D75 071101	G.S. Adams et al.	(CLEO Collab.)
ATHAR	07	PR D75 032002	S.B. Athar et al.	(CLEO Collab.)
CHEN	07B	PL B651 15	W.T. Chen et al.	(BELLE Collab.)
ABLIKIM	06D	PR D73 052006	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06I	PR D74 012004	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06R	PR D74 072001	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06T	PL B642 197	M. Ablikim et al.	(BES Collab.)
DOBBS	06	PR D73 071101	S. Dobbs et al.	(CLEO Collab.)
ABLIKIM	05G	PR D71 092002	M. Ablikim et al.	(BES Collab.)
ABLIKIM	05N	PL B630 7	M. Ablikim et al.	(BES Collab.)
ABLIKIM	05O	PL B630 21	M. Ablikim et al.	(BES Collab.)
ADAM	05A	PRL 94 232002	N.E. Adam et al.	(CLEO Collab.)
ANDREOTTI	05A	NP B717 34	M. Andreotti et al.	(FNAL E835 Collab.)
NAKAZAWA	05	PL B615 39	H. Nakazawa et al.	(BELLE Collab.)
ABLIKIM	04B	PR D70 012003	M. Ablikim et al.	(BES Collab.)
ABLIKIM	04H	PR D70 092003	M. Ablikim et al.	(BES Collab.)
ABLIKIM	04I	PR D70 092004	M. Ablikim et al.	(BES Collab.)
ATHAR	04	PR D70 112002	S.B. Athar et al.	(CLEO Collab.)
BAI	04F	PR D69 092001	J.Z. Bai et al.	(BES Collab.)
BAI	04I	PR D70 012006	J.Z. Bai et al.	(BES Collab.)
AULCHENKO	03	PL B573 63	V.M. Aukchenko et al.	(KEDR Collab.)
BAI	03C	PR D67 032004	J.Z. Bai et al.	(BES Collab.)
BAI	03E	PR D67 112001	J.Z. Bai et al.	(BES Collab.)
ABE	02T	PL B540 33	K. ABE et al.	(BELLE Collab.)
AMBROGIANI	02	PR D65 052002	M. Ambrogiani et al.	(FNAL E835 Collab.)
EISENSTEIN	01	PRL 87 061801	B.I. Eisenstein et al.	(CLEO Collab.)
AMBROGIANI	00B	PR D62 052002	M. Ambrogiani et al.	(FNAL E835 Collab.)
ACCIARRI	99E	PL B453 73	M. Acciarri et al.	(L3 Collab.)
BAI	99B	PR D60 072001	J.Z. Bai et al.	(BES Collab.)
ACKER.,K...	98	PL B439 197	K. Ackerstaff et al.	(OPAL Collab.)
BAI	98D	PR D58 092006	J.Z. Bai et al.	(BES Collab.)
BAI	98I	PRL 81 3091	J.Z. Bai et al.	(BES Collab.)
DOMINICK	94	PR D50 4265	J. Dominick et al.	(CLEO Collab.)
ARMSTRONG	93	PRL 70 2988	T.A. Armstrong et al.	(FNAL E760 Collab.)
ARMSTRONG	93E	PR D48 3037	T.A. Armstrong et al.	(FNAL E760 Collab.)
BAUER	93	PL B302 345	D.A. Bauer et al.	(TPC Collab.)
ARMSTRONG	92	NP B373 35	T.A. Armstrong et al.	(FNAL FERR, GENO+)
Also		PRL 68 1468	T.A. Armstrong et al.	(FNAL FERR, GENO+)
BAGLIN	87B	PL B187 191	C. Baglin et al.	(R704 Collab.)
BAGLIN	86B	PL B172 455	C. Baglin	(LAPP, CERN, GENO, LYON, OSLO+)
GAISER	86	PR D34 711	J. Gaiser et al.	(Crystal Ball Collab.)
LEE	85	SLAC 282	R.A. Lee	(SLAC)
LEMOIGNE	82	PL 113B 509	Y. Lemoigne et al.	(SACL, LOIC, SHMP+)
OREGLIA	82	PR D25 2259	M.J. Oreglia et al.	(SACL, CIT, HARV+)
Also		Private Comm.	M.J. Oreglia	(EFI)
BARATE	81	PR D24 2994	R. Barate et al.	(SACL, LOIC, SHMP, CERN+)
HIMEL	80	PRL 44 920	T. Himel et al.	(LBL, SLAC)
Also		Private Comm.	G. Trilling	(LBL, UCB)
BRANDELIK	79B	NP B160 426	R. Brandelik et al.	(DASP Collab.)
BARTEL	78B	PL 79B 492	W. Bartel et al.	(DESY, HEIDP)
TANENBAUM	78	PR D17 1731	W.M. Tanenbaum et al.	(SLAC, LBL)
Also		Private Comm.	G. Trilling	(LBL, UCB)
BIDDICK	77	PRL 38 1324	C.J. Biddick et al.	(UCSD, UMD, PAVI+)
WHITAKER	76	PRL 37 1596	J.S. Whitaker et al.	(SLAC, LBL)

 $\eta_c(2S)$  $I^G(J^{PC}) = 0^+(0^-+)$ 

Quantum numbers are quark model predictions.

 $\eta_c(2S)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3639.2 \pm 1.2</math> OUR AVERAGE</b>				
$3637.0 \pm 5.7 \pm 3.4$	178 <sup>1,2</sup>	LEES	14E BABR	$\gamma \gamma \rightarrow K^+ K^- \pi^0$
$3635.1 \pm 5.8 \pm 2.1$	47 <sup>1,3</sup>	LEES	14E BABR	$\gamma \gamma \rightarrow K^+ K^- \eta$
$3646.9 \pm 1.6 \pm 3.6$	57 $\pm$ 17	ABLIKIM	13k BES3	$\psi(2S) \rightarrow \gamma K_S^0 K^\pm \pi^\mp \pi^\pm \pi^\mp$
$3637.6 \pm 2.9 \pm 1.6$	127 $\pm$ 18	<sup>4</sup> ABLIKIM	12g BES3	$\psi(2S) \rightarrow \gamma K^0 K^\pm \pi^\mp \pi^\pm \pi^\mp$
$3638.5 \pm 1.5 \pm 0.8$	624	<sup>1</sup> DEL-AMO-SA...	11M BABR	$K K \pi^0 K^\pm \pi^\mp$
$3640.5 \pm 3.2 \pm 2.5$	1201	<sup>1</sup> DEL-AMO-SA...	11M BABR	$\gamma \gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$
$3636.1 \pm 3.9 \pm 0.7$ $-4.2 - 2.0$	128	<sup>5</sup> VINOKUROVA	11 BELL	$B^\pm \rightarrow K^\pm (K_S^0 K^\pm \pi^\mp)$
$3626 \pm 5 \pm 6$	311	<sup>6</sup> ABE	07 BELL	$e^+ e^- \rightarrow J/\psi(c\bar{c})$
$3645.0 \pm 5.5 \pm 4.9$ $-7.8$	121 $\pm$ 27	AUBERT	05c BABR	$e^+ e^- \rightarrow J/\psi c\bar{c}$
$3642.9 \pm 3.1 \pm 1.5$	61	ASNER	04 CLEO	$\gamma \gamma \rightarrow \eta_c \rightarrow K_S^0 K^\pm \pi^\mp$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3639 $\pm$ 7	98 $\pm$ 52	<sup>7</sup> AUBERT	06E BABR	$B^\pm \rightarrow K^\pm \chi_{c2}$
$3630.8 \pm 3.4 \pm 1.0$	112 $\pm$ 24	<sup>8</sup> AUBERT	04d BABR	$\gamma \gamma \rightarrow \eta_c(2S) \rightarrow K \bar{K} \pi$
$3654 \pm 6 \pm 8$	39 $\pm$ 11	<sup>9</sup> CHOI	02 BELL	$B \rightarrow K K_S K^- \pi^+$
$3594 \pm 5$		<sup>10</sup> EDWARDS	82c CBAL	$e^+ e^- \rightarrow \gamma X$

<sup>1</sup> Ignoring possible interference with continuum.<sup>2</sup> With a width fixed to 11.3 MeV.<sup>3</sup> With a width fixed to 11.3 MeV. Using both  $\eta \rightarrow \gamma \gamma$  and  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decays.<sup>4</sup> From a simultaneous fit to  $K_S^0 K^\pm \pi^\mp$  and  $K^+ K^- \pi^0$  decay modes.<sup>5</sup> Accounts for interference with non-resonant continuum.<sup>6</sup> From a fit of the  $J/\psi$  recoil mass spectrum. Supersedes ABE, K 02 and ABE 04g.<sup>7</sup> From the fit of the kaon momentum spectrum. Systematic errors not evaluated.<sup>8</sup> Superseded by DEL-AMO-SANCHEZ 11M.<sup>9</sup> Superseded by VINOKUROVA 11.<sup>10</sup> Assuming mass of  $\psi(2S) = 3686$  MeV.

# Meson Particle Listings

## $\eta_c(2S)$

$\eta_c(2S)$ WIDTH					
VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>11.3^{+3.2}_{-2.9}</math> OUR AVERAGE</b>					
$9.9 \pm 4.8 \pm 2.9$		$57 \pm 17$	ABLIKIM	13k BES3	$\psi(2S) \rightarrow \gamma K_S^0 K^\pm \pi^\mp \pi^\pm \pi^-$
$16.9 \pm 6.4 \pm 4.8$		$127 \pm 18$	<sup>11</sup> ABLIKIM	12G BES3	$\psi(2S) \rightarrow \gamma K^0 K \pi,$ $K K \pi^0 K^\pm \pi^\mp$
$13.4 \pm 4.6 \pm 3.2$		624	<sup>12</sup> DEL-AMO-SA...11M	BABR	$\gamma \gamma \rightarrow K_S^0 K^\pm \pi^\mp$
$6.6^{+8.4+2.6}_{-5.1-0.9}$		128	<sup>13</sup> VINOKUROVA	11 BELL	$B^\pm \rightarrow K^\pm (K_S^0 K^\pm \pi^\mp)$
$6.3 \pm 12.4 \pm 4.0$		61	ASNER	04 CLEO	$\gamma \gamma \rightarrow \eta_c \rightarrow K_S^0 K^\pm \pi^\mp$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 23	90	$98 \pm 52$	<sup>14</sup> AUBERT	06E BABR	$B^\pm \rightarrow K^\pm X_{c\overline{c}}$
$22 \pm 14$		$121 \pm 27$	AUBERT	05c BABR	$e^+ e^- \rightarrow J/\psi c \overline{c}$
$17.0 \pm 8.3 \pm 2.5$		$112 \pm 24$	<sup>15</sup> AUBERT	04D BABR	$\gamma \gamma \rightarrow \eta_c(2S) \rightarrow K \overline{K} \pi$
<55	90	$39 \pm 11$	<sup>16</sup> CHOI	02 BELL	$B \rightarrow K K_S K^- \pi^+$
<8.0	95		<sup>17</sup> EDWARDS	82c CBAL	$e^+ e^- \rightarrow \gamma X$
<sup>11</sup> From a simultaneous fit to $K_S^0 K^\pm \pi^\mp$ and $K^+ K^- \pi^0$ decay modes. <sup>12</sup> Ignoring possible interference with continuum. <sup>13</sup> Accounts for interference with non-resonant continuum. <sup>14</sup> From the fit of the kaon momentum spectrum. Systematic errors not evaluated. <sup>15</sup> Superseded by DEL-AMO-SANCHEZ 11M. <sup>16</sup> For a mass value of $3654 \pm 6$ MeV. Superseded by VINOKUROVA 11. <sup>17</sup> For a mass value of $3594 \pm 5$ MeV					

$\eta_c(2S)$ DECAY MODES			
Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level	
$\Gamma_1$ hadrons	not seen		
$\Gamma_2 K \overline{K} \pi$	( $1.9 \pm 1.2$ ) %		
$\Gamma_3 K \overline{K} \eta$	( $5 \pm 4$ ) $\times 10^{-3}$		
$\Gamma_4 2\pi^+ 2\pi^-$	not seen		
$\Gamma_5 \rho^0 \rho^0$	not seen		
$\Gamma_6 3\pi^+ 3\pi^-$	not seen		
$\Gamma_7 K^+ K^- \pi^+ \pi^-$	not seen		
$\Gamma_8 K^{*0} \overline{K}^{*0}$	not seen		
$\Gamma_9 K^+ K^- \pi^+ \pi^- \pi^0$	( $1.4 \pm 1.0$ ) %		
$\Gamma_{10} K^+ K^- 2\pi^+ 2\pi^-$	not seen		
$\Gamma_{11} K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.}$	seen		
$\Gamma_{12} 2K^+ 2K^-$	not seen		
$\Gamma_{13} \phi \phi$	not seen		
$\Gamma_{14} \rho \overline{\rho}$	< 2.0 $\times 10^{-3}$	90%	
$\Gamma_{15} \gamma \gamma$	( $1.9 \pm 1.3$ ) $\times 10^{-4}$		
$\Gamma_{16} \pi^+ \pi^- \eta$	not seen		
$\Gamma_{17} \pi^+ \pi^- \eta'$	not seen		
$\Gamma_{18} \pi^+ \pi^- \eta_c(1S)$	< 25 %	90%	

$\eta_c(2S)$ PARTIAL WIDTHS			
$\Gamma(\gamma\gamma)$	$\Gamma_{15}$		
VALUE (keV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.3 \pm 0.6$	<sup>18</sup> ASNER	04 CLEO	$\gamma \gamma \rightarrow \eta_c \rightarrow K_S^0 K^\pm \pi^\mp$
<sup>18</sup> They measure $\Gamma(\eta_c(2S) \gamma \gamma) B(\eta_c(2S) \rightarrow K \overline{K} \pi) = (0.18 \pm 0.05 \pm 0.02) \Gamma(\eta_c(1S) \gamma \gamma) B(\eta_c(1S) \rightarrow K \overline{K} \pi)$ . The value for $\Gamma(\eta_c(2S) \rightarrow \gamma \gamma)$ is derived assuming that the branching fractions for $\eta_c(2S)$ and $\eta_c(1S)$ decays to $K_S K \pi$ are equal and using $\Gamma(\eta_c(1S) \rightarrow \gamma \gamma) = 7.4 \pm 0.4 \pm 2.3$ keV.			

$\eta_c(2S) \Gamma(i) \Gamma(\gamma\gamma) / \Gamma(\text{total})$			
$\Gamma(2\pi^+ 2\pi^-) \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_4 \Gamma_{15} / \Gamma$			
VALUE (eV)	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;6.5</b>	90	UEHARA	08 BELL $\gamma \gamma \rightarrow \eta_c(2S) \rightarrow 2(\pi^+ \pi^-)$
$\Gamma(K \overline{K} \pi) \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_2 \Gamma_{15} / \Gamma$			
VALUE (eV)	EVTS	DOCUMENT ID	TECN COMMENT
<b><math>41 \pm 4 \pm 6</math></b>	624	<sup>19</sup> DEL-AMO-SA...11M	BABR $\gamma \gamma \rightarrow K_S^0 K^\pm \pi^\mp$
<sup>19</sup> Not independent from other measurements reported in DEL-AMO-SANCHEZ 11M.			
$\Gamma(K^+ K^- \pi^+ \pi^-) \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_7 \Gamma_{15} / \Gamma$			
VALUE (eV)	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;5.0</b>	90	UEHARA	08 BELL $\gamma \gamma \rightarrow \eta_c(2S) \rightarrow K^+ K^- \pi^+ \pi^-$

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_9 \Gamma_{15} / \Gamma$				
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>30 \pm 6 \pm 5</math></b>	1201	<sup>20</sup> DEL-AMO-SA...11M	BABR	$\gamma \gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$
<sup>20</sup> Not independent from other measurements reported in DEL-AMO-SANCHEZ 11M.				
$\Gamma(2K^+ 2K^-) \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_{12} \Gamma_{15} / \Gamma$				
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.9</b>	90	UEHARA	08 BELL	$\gamma \gamma \rightarrow \eta_c(2S) \rightarrow 2(K^+ K^-)$
$\Gamma(\pi^+ \pi^- \eta_c(1S)) \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_{18} \Gamma_{15} / \Gamma$				
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;133</b>	90	LEES	12AE BABR	$e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \eta_c$

$\eta_c(2S) \Gamma(i) \Gamma(\gamma\gamma) / \Gamma^2(\text{total})$				
$\Gamma(\rho \overline{\rho}) / \Gamma_{\text{total}} \times \Gamma(\gamma\gamma) / \Gamma_{\text{total}}$ $\Gamma_{14} / \Gamma \times \Gamma_{15} / \Gamma$				
VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 5.6</b>	90 <sup>21,22,23</sup>	AMBROGIANI	01 E835	$\overline{p} p \rightarrow \gamma \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 8.0	90 <sup>21,22,24</sup>	AMBROGIANI	01 E835	$\overline{p} p \rightarrow \gamma \gamma$
<12.0	90 <sup>22,24</sup>	AMBROGIANI	01 E835	$\overline{p} p \rightarrow \gamma \gamma$
<sup>21</sup> Including the measurements of of ARMSTRONG 95F in the AMBROGIANI 01 analysis. <sup>22</sup> For a total width $\Gamma=5$ MeV. <sup>23</sup> For the resonance mass region 3589–3599 MeV/ $c^2$ . <sup>24</sup> For the resonance mass region 3575–3660 MeV/ $c^2$ .				

$\eta_c(2S)$ BRANCHING RATIOS				
$\Gamma(\text{hadrons})/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma$	
VALUE		DOCUMENT ID	TECN	COMMENT
not seen		ABREU	98o DLPH	$e^+ e^- \rightarrow e^+ e^- + \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen		<sup>25</sup> EDWARDS	82c CBAL	$e^+ e^- \rightarrow \gamma X$
<sup>25</sup> For a mass value of $3594 \pm 5$ MeV				
$\Gamma(K\overline{K}\pi)/\Gamma_{\text{total}}$			$\Gamma_2/\Gamma$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.9 \pm 0.4 \pm 1.1</math></b>	$59 \pm 12$	<sup>26</sup> AUBERT	08AB BABR	$B \rightarrow \eta_c(2S) K \rightarrow K\overline{K}\pi K$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	$127 \pm 18$	ABLIKIM	13k BES3	$\psi(2S) \rightarrow \gamma K\overline{K}\pi$
seen	$39 \pm 11$	<sup>27</sup> CHOI	02 BELL	$B \rightarrow K K_S K^- \pi^+$
<sup>26</sup> Derived from a measurement of $[B(B^+ \rightarrow \eta_c(2S) K^+) \times B(\eta_c(2S) \rightarrow K\overline{K}\pi)] / [B(B^+ \rightarrow \eta_c K^+) \times B(\eta_c \rightarrow K\overline{K}\pi)] = (9.6^{+2.0}_{-1.9} \pm 2.5)\%$ and using $B(B^+ \rightarrow \eta_c(2S) K^+) = (3.4 \pm 1.8) \times 10^{-4}$ , and $[B(B^+ \rightarrow \eta_c K^+) \times B(\eta_c \rightarrow K\overline{K}\pi)] = (6.88 \pm 0.77^{+0.55}_{-0.66}) \times 10^{-5}$ .				
<sup>27</sup> For a mass value of $3654 \pm 6$ MeV				

$\Gamma(K \overline{K} \eta) / \Gamma(K \overline{K} \pi)$ $\Gamma_3 / \Gamma_2$				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>27.3 \pm 7.0 \pm 9.0</math></b>	225	<sup>28</sup> LEES	14E BABR	$\gamma \gamma \rightarrow K^+ K^- \gamma \gamma$
<sup>28</sup> LEES 14E reports $B(\eta_c(2S) \rightarrow K^+ K^- \eta) / B(\eta_c(2S) \rightarrow K^+ K^- \pi^0) = 0.82 \pm 0.21 \pm 0.27$ , which we divide by 3 to account for isospin symmetry.				

$\Gamma(2\pi^+ 2\pi^-) / \Gamma_{\text{total}}$ $\Gamma_4 / \Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>not seen</b>	UEHARA	08 BELL	$\gamma \gamma \rightarrow \eta_c(2S)$
$\Gamma(\rho^0 \rho^0) / \Gamma_{\text{total}}$ $\Gamma_5 / \Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>not seen</b>	ABLIKIM	11H BES3	$\psi(2S) \rightarrow \gamma 2\pi^+ 2\pi^-$
$\Gamma(K^+ K^- \pi^+ \pi^-) / \Gamma_{\text{total}}$ $\Gamma_7 / \Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>not seen</b>	UEHARA	08 BELL	$\gamma \gamma \rightarrow \eta_c(2S)$

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) / \Gamma(K \overline{K} \pi)$ $\Gamma_9 / \Gamma_2$				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.73 \pm 0.17 \pm 0.17</math></b>	1201	<sup>29</sup> DEL-AMO-SA...11M	BABR	$\gamma \gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$
<sup>29</sup> We have multiplied the value of $\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) / \Gamma(K_S^0 K^\pm \pi^\mp)$ reported in DEL-AMO-SANCHEZ 11M by a factor 1/3 to obtain $\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) / \Gamma(K \overline{K} \pi)$ . Not independent from other measurements reported in DEL-AMO-SANCHEZ 11M.				

$\Gamma(K^{*0} \overline{K}^{*0}) / \Gamma_{\text{total}}$ $\Gamma_8 / \Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>not seen</b>	ABLIKIM	11H BES3	$\psi(2S) \rightarrow \gamma K^+ K^- \pi^+ \pi^-$
$\Gamma(K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.}) / \Gamma_{\text{total}}$ $\Gamma_{11} / \Gamma$			
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>seen</b>	$57 \pm 17$	ABLIKIM	13K BES3 $\psi(2S) \rightarrow \gamma K_S^0 K^\pm \pi^\mp \pi^+ \pi^-$

See key on page 601

## Meson Particle Listings

 $\eta_c(2S)$ ,  $\psi(2S)$ 

$\Gamma(2K^+2K^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12}/\Gamma$
VALUE				
not seen	UEHARA	08	BELL $\gamma\gamma \rightarrow \eta_c(2S)$	

$\Gamma(\phi\phi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{13}/\Gamma$
VALUE				
not seen	ABLIKIM	11H	BES3 $\psi(2S) \rightarrow \gamma K^+ K^- K^+ K^-$	

$\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{15}/\Gamma$
VALUE					
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<5 $\times 10^{-4}$	90	<sup>30</sup> WICHT	08	BELL $B^\pm \rightarrow K^\pm \gamma\gamma$	
not seen		AMBROGIANI	01	E835 $\bar{p}p \rightarrow \gamma\gamma$	
<0.01	90	LEE	85	CBAL $\psi' \rightarrow \text{photons}$	

<sup>30</sup> WICHT 08 reports  $[\Gamma(\eta_c(2S) \rightarrow \gamma\gamma)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow \eta_c(2S) K^+)] < 0.18 \times 10^{-6}$  which we divide by our best value  $B(B^+ \rightarrow \eta_c(2S) K^+) = 3.4 \times 10^{-4}$ .

$\Gamma(\pi^+ \pi^- \eta_c(1S))/\Gamma(K\bar{K}\pi)$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{18}/\Gamma_2$
VALUE					
<3.33	90	<sup>31</sup> LEES	12AE	BABR $e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \eta_c$	

<sup>31</sup> We divided the reported limit by 3 to take into account isospin relations.

 $\eta_c(2S)$  CROSS-PARTICLE BRANCHING RATIOS

$\Gamma(\eta_c(2S) \rightarrow 2\pi^+ 2\pi^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<14.6 $\times 10^{-6}$	90	<sup>32</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma 2\pi^+ 2\pi^-$	

<sup>32</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow \rho^0 \rho^0)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<12.7 $\times 10^{-7}$	90	ABLIKIM	11H	BES3 $\psi(2S) \rightarrow \gamma 2\pi^+ 2\pi^-$	

$\Gamma(\eta_c(2S) \rightarrow 3\pi^+ 3\pi^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<13.2 $\times 10^{-6}$	90	<sup>33</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma 3\pi^+ 3\pi^-$	

<sup>33</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<9.6 $\times 10^{-6}$	90	<sup>34</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma K^+ K^- \pi^+ \pi^-$	

<sup>34</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow K^{*0} \bar{K}^{*0})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<19.6 $\times 10^{-7}$	90	ABLIKIM	11H	BES3 $\psi(2S) \rightarrow \gamma K^+ K^- \pi^+ \pi^-$	

$\Gamma(\eta_c(2S) \rightarrow K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<43.0 $\times 10^{-6}$	90	<sup>35</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma K^+ K^- \pi^+ \pi^- \pi^0$	

<sup>35</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow K^+ K^- 2\pi^+ 2\pi^-)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<9.7 $\times 10^{-6}$	90	<sup>36</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma K^+ K^- 2\pi^+ 2\pi^-$	

<sup>36</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE (units $10^{-6}$ )					
7.03 $\pm 2.10 \pm 0.7$	60	ABLIKIM	13K	BES3 $\psi(2S) \rightarrow \gamma K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.}$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$\Gamma(\eta_c(2S) \rightarrow K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE (units $10^{-6}$ )					
< 15.2	90	<sup>37</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma K_S^0 K^- 2\pi^+ \pi^- + \text{c.c.}$	

<sup>37</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow \phi\phi)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{13}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<7.8 $\times 10^{-7}$	90	ABLIKIM	11H	BES3 $\psi(2S) \rightarrow \gamma K^+ K^- K^+ K^-$	

$\Gamma(\eta_c(2S) \rightarrow \pi^+ \pi^- \eta)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{16}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<4.3 $\times 10^{-6}$	90	<sup>38</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- \eta$	

<sup>38</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow \pi^+ \pi^- \eta')/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{17}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<14.2 $\times 10^{-6}$	90	<sup>39</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- \eta'$	

<sup>39</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow K\bar{K}\eta)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<14.2 $\times 10^{-6}$	90	<sup>39</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- \eta'$	

<sup>39</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow K\bar{K}\eta)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<11.8 $\times 10^{-6}$	90	<sup>40</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma K^+ K^- \eta$	

<sup>40</sup> CRONIN-HENNESSY 10 reports a limit of  $< 5.9 \times 10^{-6}$  for the decay  $\eta_c(2S) \rightarrow K^+ K^- \eta$  which we multiply by 2 account for isospin symmetry. It assumes  $\Gamma(\eta_c(2S)) = 14$  MeV. It also gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow \pi^+ \pi^- \eta_c(1S))/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{18}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<1.7 $\times 10^{-4}$	90	<sup>41</sup> CRONIN-HEN..10	CLEO	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- \eta_c(1S)$	

<sup>41</sup> Assuming  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\eta_c(2S) \rightarrow \rho\rho)/\Gamma_{\text{total}} \times \Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{14}/\Gamma \times \Gamma_{136}^{\psi(2S)}/\Gamma\psi(2S)$
VALUE					
<1.4 $\times 10^{-6}$	90	ABLIKIM	13v	BES3 $\psi(2S) \rightarrow \gamma \rho\bar{\rho}$	

 $\eta_c(2S)$  REFERENCES

LEES	14E	PR D89 112004	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABLIKIM	13K	PR D87 052005	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	13V	PR D88 112001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	12G	PRL 109 042003	M. Ablikim <i>et al.</i>	(BES III Collab.)
LEES	12AE	PR D86 092005	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABLIKIM	11H	PR D84 091102	M. Ablikim <i>et al.</i>	(BES III Collab.)
DEL-AMO-SA...	11M	PR D84 012004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
VINOKUROVA	11	PL B706 139	A. Vinokurova <i>et al.</i>	(BELLE Collab.)
CRONIN-HEN...	10	PR D81 052002	D. Cronin-Hennessey <i>et al.</i>	(CLEO Collab.)
AUBERT	08AB	PR D78 012006	B. Aubert <i>et al.</i>	(BABAR Collab.)
UEHARA	08	EPJ C53 1	S. Uehara <i>et al.</i>	(BELLE Collab.)
WICHT	08	PL B662 323	J. Wicht <i>et al.</i>	(BELLE Collab.)
ABE	07	PRL 98 082001	K. Abe <i>et al.</i>	(BELLE Collab.)
AUBERT	06E	PRL 96 052002	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	05C	PR D72 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE	04G	PR D70 071102	K. Abe <i>et al.</i>	(BELLE Collab.)
ASNER	04	PRL 92 142001	D.M. Asner <i>et al.</i>	(CLEO Collab.)
AUBERT	04D	PR D72 142002	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE,K	02	PRL 89 142001	K. Abe <i>et al.</i>	(BELLE Collab.)
CHOI	02	PRL 89 102001	S.-K. Choi <i>et al.</i>	(BELLE Collab.)
AMBROGIANI	01	PR D64 052003	M. Ambrogiani <i>et al.</i>	(FNAL E835 Collab.)
ABREU	98O	PL B441 479	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ARMSTRONG	95F	PR D52 4839	T.A. Armstrong <i>et al.</i>	(FNAL, FERR, GENO+)
LEE	85	SLAC 282	R.A. Lee	(SLAC)
EDWARDS	82C	PRL 48 70	C. Edwards <i>et al.</i>	(CIT, HARV, PRIN+)

 $\psi(2S)$ 

$$I^G(J^{PC}) = 0^-(1^--)$$

See the Review on " $\psi(2S)$  and  $\chi_c$  branching ratios" before the  $\chi_{c0}(1P)$  Listings.

 $\psi(2S)$  MASS

OUR FIT includes measurements of  $m_{\psi(2S)}$ ,  $m_{\psi(3770)}$ , and  $m_{\psi(3770)} - m_{\psi(2S)}$ .

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
3686.097 $\pm 0.025$ OUR FIT	Error	includes scale factor of 2.6.		
3686.097 $\pm 0.010$ OUR AVERAGE				
3686.099 $\pm 0.004 \pm 0.009$		<sup>1</sup> ANASHIN	15	KEDR $e^+ e^- \rightarrow \text{hadrons}$
3686.12 $\pm 0.06 \pm 0.10$	4k	AAIJ	12H	LHCB $p\bar{p} \rightarrow J/\psi \pi^+ \pi^- X$
3685.95 $\pm 0.10$	413	<sup>2</sup> ARTAMONOV	00	OLYA $e^+ e^- \rightarrow \text{hadrons}$
3685.98 $\pm 0.09 \pm 0.04$		<sup>3</sup> ARMSTRONG	93b	E760 $\bar{p}p \rightarrow e^+ e^-$

# Meson Particle Listings

## $\psi(2S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3686.114 \pm 0.007^{+0.011}_{-0.016}$	<sup>4</sup> ANASHIN	12	KEDR	$e^+e^- \rightarrow \text{hadrons}$
$3686.111 \pm 0.025 \pm 0.009$	AULCHENKO	03	KEDR	$e^+e^- \rightarrow \text{hadrons}$
$3686.00 \pm 0.10$	<sup>413</sup> <sup>5</sup> ZHOLENTZ	80	OLYA	$e^+e^-$
<sup>1</sup> Supersedes AULCHENKO 03 and ANASHIN 12.				
<sup>2</sup> Reanalysis of ZHOLENTZ 80 using new electron mass (COHEN 87) and radiative corrections (KURAEV 85).				
<sup>3</sup> Mass central value and systematic error recalculated by us according to Eq.(16) in ARMSTRONG 93B, using the value for the $J/\psi(1S)$ mass from AULCHENKO 03.				
<sup>4</sup> From the scans in 2004 and 2006. ANASHIN 12 reports the value $3686.114 \pm 0.007 \pm 0.011^{+0.002}_{-0.012}$ MeV, where the third uncertainty is due to assumptions on the interference between the resonance and hadronic continuum. We combined the two systematic uncertainties.				
<sup>5</sup> Superseded by ARTAMONOV 00.				

### $m_{\psi(2S)} - m_{J/\psi(1S)}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>589.188 ± 0.028 OUR AVERAGE</b>			
$589.194 \pm 0.027 \pm 0.011$	<sup>1</sup> AULCHENKO 03	KEDR	$e^+e^- \rightarrow \text{hadrons}$
$589.7 \pm 1.2$	LEMOIGNE 82	GOLI	$185 \pi^- \text{Be} \rightarrow \gamma \mu^+ \mu^- \text{A}$
$589.07 \pm 0.13$	<sup>1</sup> ZHOLENTZ 80	OLYA	$e^+e^-$
$588.7 \pm 0.8$	LUTH 75	MRK1	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$588 \pm 1$	<sup>2</sup> BAI	98E BES	$e^+e^-$
<sup>1</sup> Redundant with data in mass above.			
<sup>2</sup> Systematic errors not evaluated.			

### $\psi(2S)$ WIDTH

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>296 ± 8 OUR FIT</b>				
<b>286 ± 16 OUR AVERAGE</b>				
$358 \pm 88 \pm 4$		ABLIKIM 08B	BES2	$e^+e^- \rightarrow \text{hadrons}$
$290 \pm 25 \pm 4$	2.7k	ANDREOTTI 07	E835	$p\bar{p} \rightarrow e^+e^-, J/\psi X$
$331 \pm 58 \pm 2$		ABLIKIM 06L	BES2	$e^+e^- \rightarrow \text{hadrons}$
$264 \pm 27$		<sup>1</sup> BAI 02B	BES2	$e^+e^-$
$287 \pm 37 \pm 16$		<sup>2</sup> ARMSTRONG 93B	E760	$\bar{p}p \rightarrow e^+e^-$
<sup>1</sup> From a simultaneous fit to the hadronic and $\mu^+\mu^-$ cross section, assuming $\Gamma = \Gamma_h + \Gamma_e + \Gamma_\mu + \Gamma_\tau$ and lepton universality. Does not include vacuum polarization correction.				
<sup>2</sup> The initial-state radiation correction reevaluated by ANDREOTTI 07 in its Ref. [4].				

### $\psi(2S)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ hadrons	$(97.85 \pm 0.13) \%$	
$\Gamma_2$ virtual $\gamma \rightarrow \text{hadrons}$	$(1.73 \pm 0.14) \%$	S=1.5
$\Gamma_3$ $ggg$	$(10.6 \pm 1.6) \%$	
$\Gamma_4$ $\gamma g g$	$(1.03 \pm 0.29) \%$	
$\Gamma_5$ light hadrons	$(15.4 \pm 1.5) \%$	
$\Gamma_6$ $e^+e^-$	$(7.89 \pm 0.17) \times 10^{-3}$	
$\Gamma_7$ $\mu^+\mu^-$	$(7.9 \pm 0.9) \times 10^{-3}$	
$\Gamma_8$ $\tau^+\tau^-$	$(3.1 \pm 0.4) \times 10^{-3}$	

### Decays into $J/\psi(1S)$ and anything

$\Gamma_9$ $J/\psi(1S)$ anything	$(61.0 \pm 0.6) \%$	
$\Gamma_{10}$ $J/\psi(1S)$ neutrals	$(25.14 \pm 0.33) \%$	
$\Gamma_{11}$ $J/\psi(1S) \pi^+\pi^-$	$(34.49 \pm 0.30) \%$	
$\Gamma_{12}$ $J/\psi(1S) \pi^0 \pi^0$	$(18.16 \pm 0.31) \%$	
$\Gamma_{13}$ $J/\psi(1S) \eta$	$(3.36 \pm 0.05) \%$	
$\Gamma_{14}$ $J/\psi(1S) \pi^0$	$(1.268 \pm 0.032) \times 10^{-3}$	

### Hadronic decays

$\Gamma_{15}$ $\pi^0 h_c(1P)$	$(8.6 \pm 1.3) \times 10^{-4}$	
$\Gamma_{16}$ $3(\pi^+\pi^-)\pi^0$	$(3.5 \pm 1.6) \times 10^{-3}$	
$\Gamma_{17}$ $2(\pi^+\pi^-)\pi^0$	$(2.9 \pm 1.0) \times 10^{-3}$	S=4.7
$\Gamma_{18}$ $\rho \partial_2(1320)$	$(2.6 \pm 0.9) \times 10^{-4}$	
$\Gamma_{19}$ $\rho \bar{p}$	$(2.88 \pm 0.09) \times 10^{-4}$	
$\Gamma_{20}$ $\Delta^{++} \bar{\Delta}^{--}$	$(1.28 \pm 0.35) \times 10^{-4}$	
$\Gamma_{21}$ $\Lambda \bar{\Lambda} \pi^0$	$< 2.9 \times 10^{-6}$	CL=90%
$\Gamma_{22}$ $\Lambda \bar{\Lambda} \eta$	$(2.5 \pm 0.4) \times 10^{-5}$	
$\Gamma_{23}$ $\Lambda \bar{p} K^+$	$(1.00 \pm 0.14) \times 10^{-4}$	
$\Gamma_{24}$ $\Lambda \bar{p} K^+ \pi^+ \pi^-$	$(1.8 \pm 0.4) \times 10^{-4}$	
$\Gamma_{25}$ $\Lambda \bar{\Lambda} \pi^+ \pi^-$	$(2.8 \pm 0.6) \times 10^{-4}$	
$\Gamma_{26}$ $\Lambda \bar{\Lambda}$	$(3.57 \pm 0.18) \times 10^{-4}$	
$\Gamma_{27}$ $\Lambda \bar{\Sigma}^+ \pi^- + \text{c.c.}$	$(1.40 \pm 0.13) \times 10^{-4}$	
$\Gamma_{28}$ $\Lambda \bar{\Sigma}^- \pi^+ + \text{c.c.}$	$(1.54 \pm 0.14) \times 10^{-4}$	
$\Gamma_{29}$ $\Sigma^0 \bar{p} K^+ + \text{c.c.}$	$(1.67 \pm 0.18) \times 10^{-5}$	
$\Gamma_{30}$ $\Sigma^+ \bar{\Sigma}^-$	$(2.51 \pm 0.21) \times 10^{-4}$	

$\Gamma_{31}$ $\Sigma^0 \bar{\Sigma}^0$	$(2.32 \pm 0.16) \times 10^{-4}$	
$\Gamma_{32}$ $\Sigma(1385)^+ \bar{\Sigma}(1385)^-$	$(1.1 \pm 0.4) \times 10^{-4}$	
$\Gamma_{33}$ $\Xi^- \bar{\Xi}^+$	$(2.64 \pm 0.18) \times 10^{-4}$	
$\Gamma_{34}$ $\Xi^0 \bar{\Xi}^0$	$(2.07 \pm 0.23) \times 10^{-4}$	
$\Gamma_{35}$ $\Xi(1530)^0 \bar{\Xi}(1530)^0$	$(5.2 \pm^{+3.2}_{-1.2}) \times 10^{-5}$	
$\Gamma_{36}$ $K^- \Lambda \Xi^+ + \text{c.c.}$	$(3.9 \pm 0.4) \times 10^{-5}$	
$\Gamma_{37}$ $\Xi(1690)^- \bar{\Xi}^+ \rightarrow K^- \Lambda \Xi^+ + \text{c.c.}$	$(5.2 \pm 1.6) \times 10^{-6}$	
$\Gamma_{38}$ $\Xi(1820)^- \bar{\Xi}^+ \rightarrow K^- \Lambda \Xi^+ + \text{c.c.}$	$(1.20 \pm 0.32) \times 10^{-5}$	
$\Gamma_{39}$ $K^- \Sigma^0 \Xi^+ + \text{c.c.}$	$(3.7 \pm 0.4) \times 10^{-5}$	
$\Gamma_{40}$ $\Omega^- \bar{\Omega}^+$	$(4.7 \pm 1.0) \times 10^{-5}$	
$\Gamma_{41}$ $\pi^0 \rho \bar{p}$	$(1.53 \pm 0.07) \times 10^{-4}$	
$\Gamma_{42}$ $N(940) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(6.4 \pm^{+1.8}_{-1.3}) \times 10^{-5}$	
$\Gamma_{43}$ $N(1440) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(7.3 \pm^{+1.7}_{-1.5}) \times 10^{-5}$	S=2.5
$\Gamma_{44}$ $N(1520) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(6.4 \pm^{+2.3}_{-1.8}) \times 10^{-6}$	
$\Gamma_{45}$ $N(1535) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(2.5 \pm 1.0) \times 10^{-5}$	
$\Gamma_{46}$ $N(1650) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(3.8 \pm^{+1.4}_{-1.7}) \times 10^{-5}$	
$\Gamma_{47}$ $N(1720) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(1.79 \pm^{+0.26}_{-0.70}) \times 10^{-5}$	
$\Gamma_{48}$ $N(2300) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(2.6 \pm^{+1.2}_{-0.7}) \times 10^{-5}$	
$\Gamma_{49}$ $N(2570) \bar{p} + \text{c.c.} \rightarrow \pi^0 \rho \bar{p}$	$(2.13 \pm^{+0.40}_{-0.31}) \times 10^{-5}$	
$\Gamma_{50}$ $\pi^0 f_0(2100) \rightarrow \pi^0 \rho \bar{p}$	$(1.1 \pm 0.4) \times 10^{-5}$	
$\Gamma_{51}$ $\eta \rho \bar{p}$	$(6.0 \pm 0.4) \times 10^{-5}$	
$\Gamma_{52}$ $\eta f_0(2100) \rightarrow \eta \rho \bar{p}$	$(1.2 \pm 0.4) \times 10^{-5}$	
$\Gamma_{53}$ $N(1535) \bar{p} \rightarrow \eta \rho \bar{p}$	$(4.4 \pm 0.7) \times 10^{-5}$	
$\Gamma_{54}$ $\omega \rho \bar{p}$	$(6.9 \pm 2.1) \times 10^{-5}$	
$\Gamma_{55}$ $\phi \rho \bar{p}$	$< 2.4 \times 10^{-5}$	CL=90%
$\Gamma_{56}$ $\pi^+ \pi^- \rho \bar{p}$	$(6.0 \pm 0.4) \times 10^{-4}$	
$\Gamma_{57}$ $\rho \bar{\pi} \pi^-$ or c.c.	$(2.48 \pm 0.17) \times 10^{-4}$	
$\Gamma_{58}$ $\rho \bar{\pi} \pi^- \pi^0$	$(3.2 \pm 0.7) \times 10^{-4}$	
$\Gamma_{59}$ $2(\pi^+ \pi^- \pi^0)$	$(4.8 \pm 1.5) \times 10^{-3}$	
$\Gamma_{60}$ $\eta \pi^+ \pi^-$	$< 1.6 \times 10^{-4}$	CL=90%
$\Gamma_{61}$ $\eta \pi^+ \pi^- \pi^0$	$(9.5 \pm 1.7) \times 10^{-4}$	
$\Gamma_{62}$ $2(\pi^+ \pi^-) \eta$	$(1.2 \pm 0.6) \times 10^{-3}$	
$\Gamma_{63}$ $\eta' \pi^+ \pi^- \pi^0$	$(4.5 \pm 2.1) \times 10^{-4}$	
$\Gamma_{64}$ $\omega \pi^+ \pi^-$	$(7.3 \pm 1.2) \times 10^{-4}$	S=2.1
$\Gamma_{65}$ $b_1^\pm \pi^\mp$	$(4.0 \pm 0.6) \times 10^{-4}$	S=1.1
$\Gamma_{66}$ $b_1^0 \pi^0$	$(2.4 \pm 0.6) \times 10^{-4}$	
$\Gamma_{67}$ $\omega f_2(1270)$	$(2.2 \pm 0.4) \times 10^{-4}$	
$\Gamma_{68}$ $\pi^+ \pi^- K^+ K^-$	$(7.5 \pm 0.9) \times 10^{-4}$	S=1.9
$\Gamma_{69}$ $\rho^0 K^+ K^-$	$(2.2 \pm 0.4) \times 10^{-4}$	
$\Gamma_{70}$ $K^*(892)^0 \bar{K}_2^*(1430)^0$	$(1.9 \pm 0.5) \times 10^{-4}$	
$\Gamma_{71}$ $K^+ K^- \pi^+ \pi^- \eta$	$(1.3 \pm 0.7) \times 10^{-3}$	
$\Gamma_{72}$ $K^+ K^- 2(\pi^+ \pi^-) \pi^0$	$(1.00 \pm 0.31) \times 10^{-3}$	
$\Gamma_{73}$ $K^+ K^- 2(\pi^+ \pi^-)$	$(1.9 \pm 0.9) \times 10^{-3}$	
$\Gamma_{74}$ $K_1(1270)^\pm K^\mp$	$(1.00 \pm 0.28) \times 10^{-3}$	
$\Gamma_{75}$ $K_S^0 K_S^0 \pi^+ \pi^-$	$(2.2 \pm 0.4) \times 10^{-4}$	
$\Gamma_{76}$ $\rho^0 \rho \bar{p}$	$(5.0 \pm 2.2) \times 10^{-5}$	
$\Gamma_{77}$ $K^+ \bar{K}^*(892)^0 \pi^- + \text{c.c.}$	$(6.7 \pm 2.5) \times 10^{-4}$	
$\Gamma_{78}$ $2(\pi^+ \pi^-)$	$(2.4 \pm 0.6) \times 10^{-4}$	S=2.2
$\Gamma_{79}$ $\rho^0 \pi^+ \pi^-$	$(2.2 \pm 0.6) \times 10^{-4}$	S=1.4
$\Gamma_{80}$ $K^+ K^- \pi^+ \pi^- \pi^0$	$(1.26 \pm 0.09) \times 10^{-3}$	
$\Gamma_{81}$ $\omega f_0(1710) \rightarrow \omega K^+ K^-$	$(5.9 \pm 2.2) \times 10^{-5}$	
$\Gamma_{82}$ $K^*(892)^0 K^- \pi^+ \pi^0 + \text{c.c.}$	$(8.6 \pm 2.2) \times 10^{-4}$	
$\Gamma_{83}$ $K^*(892)^+ K^- \pi^+ \pi^- + \text{c.c.}$	$(9.6 \pm 2.8) \times 10^{-4}$	
$\Gamma_{84}$ $K^*(892)^+ K^- \rho^0 + \text{c.c.}$	$(7.3 \pm 2.6) \times 10^{-4}$	
$\Gamma_{85}$ $K^*(892)^0 K^- \rho^+ + \text{c.c.}$	$(6.1 \pm 1.8) \times 10^{-4}$	
$\Gamma_{86}$ $\eta K^+ K^-$ , no $\eta \phi$	$(3.1 \pm 0.4) \times 10^{-5}$	
$\Gamma_{87}$ $\omega K^+ K^-$	$(1.62 \pm 0.11) \times 10^{-4}$	S=1.1
$\Gamma_{88}$ $\omega K^*(892)^+ K^- + \text{c.c.}$	$(2.07 \pm 0.26) \times 10^{-4}$	
$\Gamma_{89}$ $\omega K_2^*(1430)^+ K^- + \text{c.c.}$	$(6.1 \pm 1.2) \times 10^{-5}$	
$\Gamma_{90}$ $\omega \bar{K}^*(892)^0 K^0$	$(1.68 \pm 0.30) \times 10^{-4}$	
$\Gamma_{91}$ $\omega \bar{K}_2^*(1430)^0 K^0$	$(5.8 \pm 2.2) \times 10^{-5}$	
$\Gamma_{92}$ $\omega X(1440) \rightarrow \omega K_S^0 K^- \pi^+ + \text{c.c.}$	$(1.6 \pm 0.4) \times 10^{-5}$	
$\Gamma_{93}$ $\omega X(1440) \rightarrow \omega K^+ K^- \pi^0$	$(1.09 \pm 0.26) \times 10^{-5}$	
$\Gamma_{94}$ $\omega f_1(1285) \rightarrow \omega K_S^0 K^- \pi^+ + \text{c.c.}$	$(3.0 \pm 1.0) \times 10^{-6}$	
$\Gamma_{95}$ $\omega f_1(1285) \rightarrow \omega K^+ K^- \pi^0$	$(1.2 \pm 0.7) \times 10^{-6}$	
$\Gamma_{96}$ $3(\pi^+ \pi^-)$	$(3.5 \pm 2.0) \times 10^{-4}$	S=2.8
$\Gamma_{97}$ $\rho \bar{p} \pi^+ \pi^- \pi^0$	$(7.3 \pm 0.7) \times 10^{-4}$	

$\Gamma_{98}$	$K^+ K^-$	$(7.5 \pm 0.5) \times 10^{-5}$	
$\Gamma_{99}$	$K_S^0 \bar{K}_L^0$	$(5.34 \pm 0.33) \times 10^{-5}$	
$\Gamma_{100}$	$\pi^+ \pi^- \pi^0$	$(2.01 \pm 0.17) \times 10^{-4}$	S=1.7
$\Gamma_{101}$	$\rho(2150) \pi \rightarrow \pi^+ \pi^- \pi^0$	$(1.9 \pm 1.2) \times 10^{-4}$	
$\Gamma_{102}$	$\rho(770) \pi \rightarrow \pi^+ \pi^- \pi^0$	$(3.2 \pm 1.2) \times 10^{-5}$	S=1.8
$\Gamma_{103}$	$\pi^+ \pi^-$	$(7.8 \pm 2.6) \times 10^{-6}$	
$\Gamma_{104}$	$K_1(1400)^\pm K^\mp$	$< 3.1 \times 10^{-4}$	CL=90%
$\Gamma_{105}$	$K_2^*(1430)^\pm K^\mp$	$(7.1 \pm 1.3) \times 10^{-5}$	
$\Gamma_{106}$	$K^+ K^- \pi^0$	$(4.07 \pm 0.31) \times 10^{-5}$	
$\Gamma_{107}$	$K^+ K^*(892)^- + \text{c.c.}$	$(2.9 \pm 0.4) \times 10^{-5}$	S=1.2
$\Gamma_{108}$	$K^*(892)^0 \bar{K}^0 + \text{c.c.}$	$(1.09 \pm 0.20) \times 10^{-4}$	
$\Gamma_{109}$	$\phi \pi^+ \pi^-$	$(1.17 \pm 0.29) \times 10^{-4}$	S=1.7
$\Gamma_{110}$	$\phi f_0(980) \rightarrow \pi^+ \pi^-$	$(6.8 \pm 2.5) \times 10^{-5}$	S=1.2
$\Gamma_{111}$	$2(K^+ K^-)$	$(6.0 \pm 1.4) \times 10^{-5}$	
$\Gamma_{112}$	$\phi K^+ K^-$	$(7.0 \pm 1.6) \times 10^{-5}$	
$\Gamma_{113}$	$2(K^+ K^-) \pi^0$	$(1.10 \pm 0.28) \times 10^{-4}$	
$\Gamma_{114}$	$\phi \eta$	$(3.10 \pm 0.31) \times 10^{-5}$	
$\Gamma_{115}$	$\phi \eta'$	$(3.1 \pm 1.6) \times 10^{-5}$	
$\Gamma_{116}$	$\omega \eta'$	$(3.2 \pm 2.5) \times 10^{-5}$	
$\Gamma_{117}$	$\omega \pi^0$	$(2.1 \pm 0.6) \times 10^{-5}$	
$\Gamma_{118}$	$\rho \eta'$	$(1.9 \pm 1.7) \times 10^{-5}$	
$\Gamma_{119}$	$\rho \eta$	$(2.2 \pm 0.6) \times 10^{-5}$	S=1.1
$\Gamma_{120}$	$\omega \eta$	$< 1.1 \times 10^{-5}$	CL=90%
$\Gamma_{121}$	$\phi \pi^0$	$< 4 \times 10^{-7}$	CL=90%
$\Gamma_{122}$	$\eta_c \pi^+ \pi^- \pi^0$	$< 1.0 \times 10^{-3}$	CL=90%
$\Gamma_{123}$	$p \bar{p} K^+ K^-$	$(2.7 \pm 0.7) \times 10^{-5}$	
$\Gamma_{124}$	$\bar{\Lambda} n K_S^0 + \text{c.c.}$	$(8.1 \pm 1.8) \times 10^{-5}$	
$\Gamma_{125}$	$\phi f_2'(1525)$	$(4.4 \pm 1.6) \times 10^{-5}$	
$\Gamma_{126}$	$\Theta(1540) \bar{\Theta}(1540) \rightarrow K_S^0 p K^- \bar{\pi} + \text{c.c.}$	$< 8.8 \times 10^{-6}$	CL=90%
$\Gamma_{127}$	$\Theta(1540) K^- \bar{\pi} \rightarrow K_S^0 p K^- \bar{\pi}$	$< 1.0 \times 10^{-5}$	CL=90%
$\Gamma_{128}$	$\Theta(1540) K_S^0 \bar{p} \rightarrow K_S^0 \bar{p} K^+ n$	$< 7.0 \times 10^{-6}$	CL=90%
$\Gamma_{129}$	$\bar{\Theta}(1540) K^+ n \rightarrow K_S^0 \bar{p} K^+ n$	$< 2.6 \times 10^{-5}$	CL=90%
$\Gamma_{130}$	$\bar{\Theta}(1540) K_S^0 p \rightarrow K_S^0 p K^- \bar{\pi}$	$< 6.0 \times 10^{-6}$	CL=90%
$\Gamma_{131}$	$K_S^0 \bar{K}_S^0$	$< 4.6 \times 10^{-6}$	

## Radiative decays

$\Gamma_{132}$	$\gamma \chi_{c0}(1P)$	$(9.99 \pm 0.27) \%$	
$\Gamma_{133}$	$\gamma \chi_{c1}(1P)$	$(9.55 \pm 0.31) \%$	
$\Gamma_{134}$	$\gamma \chi_{c2}(1P)$	$(9.11 \pm 0.31) \%$	
$\Gamma_{135}$	$\gamma \eta_c(1S)$	$(3.4 \pm 0.5) \times 10^{-3}$	S=1.3
$\Gamma_{136}$	$\gamma \eta_c(2S)$	$(7 \pm 5) \times 10^{-4}$	
$\Gamma_{137}$	$\gamma \pi^0$	$(1.6 \pm 0.4) \times 10^{-6}$	
$\Gamma_{138}$	$\gamma \eta'(958)$	$(1.23 \pm 0.06) \times 10^{-4}$	
$\Gamma_{139}$	$\gamma f_2(1270)$	$(2.73 \pm 0.29) \times 10^{-4}$	S=1.8
$\Gamma_{140}$	$\gamma f_0(1370) \rightarrow \gamma K \bar{K}$	$(3.1 \pm 1.7) \times 10^{-5}$	
$\Gamma_{141}$	$\gamma f_0(1500)$	$(9.2 \pm 1.9) \times 10^{-5}$	
$\Gamma_{142}$	$\gamma f_2'(1525)$	$(3.3 \pm 0.8) \times 10^{-5}$	
$\Gamma_{143}$	$\gamma f_0(1710)$		
$\Gamma_{144}$	$\gamma f_0(1710) \rightarrow \gamma \pi \pi$	$(3.5 \pm 0.6) \times 10^{-5}$	
$\Gamma_{145}$	$\gamma f_0(1710) \rightarrow \gamma K \bar{K}$	$(6.6 \pm 0.7) \times 10^{-5}$	
$\Gamma_{146}$	$\gamma f_0(2100) \rightarrow \gamma \pi \pi$	$(4.8 \pm 1.0) \times 10^{-6}$	
$\Gamma_{147}$	$\gamma f_0(2200) \rightarrow \gamma K \bar{K}$	$(3.2 \pm 1.0) \times 10^{-6}$	
$\Gamma_{148}$	$\gamma f_J(2220) \rightarrow \gamma \pi \pi$	$< 5.8 \times 10^{-6}$	CL=90%
$\Gamma_{149}$	$\gamma f_J(2220) \rightarrow \gamma K \bar{K}$	$< 9.5 \times 10^{-6}$	CL=90%
$\Gamma_{150}$	$\gamma \gamma$	$< 1.5 \times 10^{-4}$	CL=90%
$\Gamma_{151}$	$\gamma \eta$	$(1.4 \pm 0.5) \times 10^{-6}$	
$\Gamma_{152}$	$\gamma \eta \pi^+ \pi^-$	$(8.7 \pm 2.1) \times 10^{-4}$	
$\Gamma_{153}$	$\gamma \eta(1405)$		
$\Gamma_{154}$	$\gamma \eta(1405) \rightarrow \gamma K \bar{K} \pi$	$< 9 \times 10^{-5}$	CL=90%
$\Gamma_{155}$	$\gamma \eta(1405) \rightarrow \eta \pi^+ \pi^-$	$(3.6 \pm 2.5) \times 10^{-5}$	
$\Gamma_{156}$	$\gamma \eta(1475)$		
$\Gamma_{157}$	$\gamma \eta(1475) \rightarrow K \bar{K} \pi$	$< 1.4 \times 10^{-4}$	CL=90%
$\Gamma_{158}$	$\gamma \eta(1475) \rightarrow \eta \pi^+ \pi^-$	$< 8.8 \times 10^{-5}$	CL=90%
$\Gamma_{159}$	$\gamma 2(\pi^+ \pi^-)$	$(4.0 \pm 0.6) \times 10^{-4}$	
$\Gamma_{160}$	$\gamma K^{*0} K^+ \pi^- + \text{c.c.}$	$(3.7 \pm 0.9) \times 10^{-4}$	
$\Gamma_{161}$	$\gamma K^{*0} \bar{K}^{*0}$	$(2.4 \pm 0.7) \times 10^{-4}$	
$\Gamma_{162}$	$\gamma K_S^0 K^+ \pi^- + \text{c.c.}$	$(2.6 \pm 0.5) \times 10^{-4}$	
$\Gamma_{163}$	$\gamma K^+ K^- \pi^+ \pi^-$	$(1.9 \pm 0.5) \times 10^{-4}$	
$\Gamma_{164}$	$\gamma p \bar{p}$	$(3.9 \pm 0.5) \times 10^{-5}$	S=2.0
$\Gamma_{165}$	$\gamma f_2(1950) \rightarrow \gamma p \bar{p}$	$(1.20 \pm 0.22) \times 10^{-5}$	
$\Gamma_{166}$	$\gamma f_2(2150) \rightarrow \gamma p \bar{p}$	$(7.2 \pm 1.8) \times 10^{-6}$	

$\Gamma_{167}$	$\gamma X(1835) \rightarrow \gamma p \bar{p}$	$(4.6 \pm 1.8) \times 10^{-6}$	
$\Gamma_{168}$	$\gamma X \rightarrow \gamma p \bar{p}$	$[a] < 2 \times 10^{-6}$	CL=90%
$\Gamma_{169}$	$\gamma \pi^+ \pi^- p \bar{p}$	$(2.8 \pm 1.4) \times 10^{-5}$	
$\Gamma_{170}$	$\gamma 2(\pi^+ \pi^-) K^+ K^-$	$< 2.2 \times 10^{-4}$	CL=90%
$\Gamma_{171}$	$\gamma 3(\pi^+ \pi^-)$	$< 1.7 \times 10^{-4}$	CL=90%
$\Gamma_{172}$	$\gamma K^+ K^- K^+ K^-$	$< 4 \times 10^{-5}$	CL=90%
$\Gamma_{173}$	$\gamma \gamma J/\psi$	$(3.1 \pm 1.0) \times 10^{-4}$	

## Other decays

$\Gamma_{174}$	invisible	$< 1.6 \%$	CL=90%
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[a] For a narrow resonance in the range  $2.2 < M(X) < 2.8$  GeV.

## CONSTRAINED FIT INFORMATION

A multiparticle fit to  $\chi_{c1}(1P)$ ,  $\chi_{c0}(1P)$ ,  $\chi_{c2}(1P)$ , and  $\psi(2S)$  with 4 total widths, a partial width, 25 combinations of partial widths obtained from integrated cross section, and 84 branching ratios uses 240 measurements to determine 49 parameters. The overall fit has a  $\chi^2 = 342.4$  for 191 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ .

$x_7$	3								
$x_8$	1	0							
$x_{11}$	30	8	2						
$x_{12}$	29	5	1	49					
$x_{13}$	13	3	1	36	16				
$x_{19}$	0	0	0	5	3	2			
$x_{132}$	1	0	0	3	1	1	0		
$x_{133}$	2	0	0	4	1	1	0	0	
$x_{134}$	1	0	0	4	1	1	0	0	0
$\Gamma$	-81	-3	-1	-39	-35	-17	-9	-1	-2
	$x_6$	$x_7$	$x_8$	$x_{11}$	$x_{12}$	$x_{13}$	$x_{19}$	$x_{132}$	$x_{133}$

 $\psi(2S)$  PARTIAL WIDTHS

$\Gamma(\text{hadrons})$				$\Gamma_1$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
258 ± 26	BAI	02B	BES2 $e^+ e^-$	
224 ± 56	LUTH	75	MRK1 $e^+ e^-$	
$\Gamma(e^+ e^-)$				$\Gamma_6$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
<b>2.34 ± 0.04 OUR FIT</b>				
<b>2.30 ± 0.06 OUR AVERAGE</b>				
2.24 ± 0.10 ± 0.02	<sup>1</sup> ABLIKIM	15v	BES3 $4.0\text{--}4.4 e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
2.338 ± 0.037 ± 0.096	ABLIKIM	08B	BES2 $e^+ e^- \rightarrow \text{hadrons}$	
2.330 ± 0.036 ± 0.110	ABLIKIM	06L	BES2 $e^+ e^- \rightarrow \text{hadrons}$	
2.44 ± 0.21	<sup>2</sup> BAI	02B	BES2 $e^+ e^-$	
2.14 ± 0.21	ALEXANDER	89	RVUE See T mini-review	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.0 ± 0.3	BRANDELIK	79c	DASP $e^+ e^-$	
2.1 ± 0.3	<sup>3</sup> LUTH	75	MRK1 $e^+ e^-$	
<sup>1</sup> ABLIKIM 15v reports $2.213 \pm 0.018 \pm 0.099$ keV from a measurement of $[\Gamma(\psi(2S) \rightarrow e^+ e^-)] \times [B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)]$ assuming $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (34.95 \pm 0.45) \times 10^{-2}$ , which we rescale to our best value $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (34.49 \pm 0.30) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>2</sup> From a simultaneous fit to $e^+ e^-$ , $\mu^+ \mu^-$ , and hadronic channel, assuming $\Gamma_e = \Gamma_\mu = \Gamma_\gamma / 0.38847$ .				
<sup>3</sup> From a simultaneous fit to $e^+ e^-$ , $\mu^+ \mu^-$ , and hadronic channels assuming $\Gamma(e^+ e^-) = \Gamma(\mu^+ \mu^-)$ .				
$\Gamma(\gamma\gamma)$				$\Gamma_{150}$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 43</b>	90	BRANDELIK	79c	DASP $e^+ e^-$

# Meson Particle Listings

## $\psi(2S)$

### $\psi(2S) \Gamma(i) \Gamma(e^+ e^-) / \Gamma(\text{total})$

This combination of a partial width with the partial width into  $e^+ e^-$  and with the total width is obtained from the integrated cross section into channel(i) in the  $e^+ e^-$  annihilation. We list only data that have not been used to determine the partial width  $\Gamma(i)$  or the branching ratio  $\Gamma(i)/\text{total}$ .

### $\Gamma(\text{hadrons}) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_1 \Gamma_6 / \Gamma$
<b>2.233 ± 0.015 ± 0.042</b>	<sup>1</sup> ANASHIN	12	KEDR $e^+ e^- \rightarrow \text{hadrons}$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

2.2 ± 0.4	ABRAMS	75	MRK1 $e^+ e^-$	
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<sup>1</sup> ANASHIN 12 reports the value  $2.233 \pm 0.015 \pm 0.037 \pm 0.020$  keV, where the third uncertainty is due to assumptions on the interference between the resonance and hadronic continuum. We combined the two systematic uncertainties.

### $\Gamma(\tau^+ \tau^-) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_8 \Gamma_6 / \Gamma$
9.0 ± 2.6	79	<sup>1</sup> ANASHIN	07	KEDR $e^+ e^- \rightarrow \psi(2S) \rightarrow \tau^+ \tau^-$	

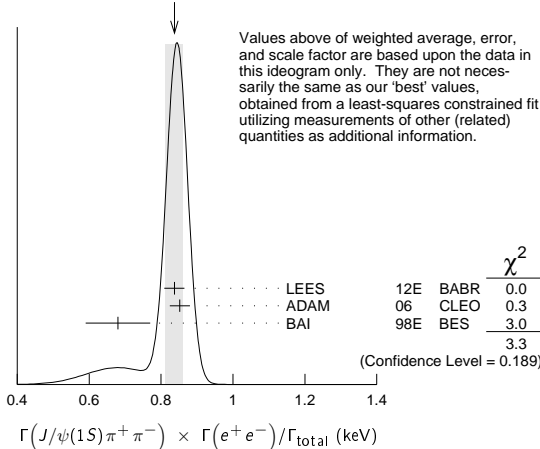
<sup>1</sup> Using  $\psi(2S)$  total width of  $337 \pm 13$  keV. Systematic errors not evaluated.

### $\Gamma(J/\psi(1S) \pi^+ \pi^-) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11} \Gamma_6 / \Gamma$
<b>0.807 ± 0.013 OUR FIT</b>					
<b>0.837 ± 0.025 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.	
0.837 ± 0.028 ± 0.005		<sup>1</sup> LEES	12E	BABR $10.6 e^+ e^- \rightarrow 2\pi^+ 2\pi^- \gamma$	
0.852 ± 0.010 ± 0.026	19.5k	ADAM	06	CLEO $3.773 e^+ e^- \rightarrow \gamma \psi(2S)$	
0.68 ± 0.09		<sup>2</sup> BAI	98E	BES $e^+ e^-$	
0.88 ± 0.08 ± 0.03	256	<sup>3</sup> AUBERT	07AU	BABR $10.6 e^+ e^- \rightarrow J/\psi \pi^+ \pi^- \gamma$	
0.755 ± 0.048 ± 0.004	544	<sup>4</sup> AUBERT	05D	BABR $10.6 e^+ e^- \rightarrow \pi^+ \pi^- \mu^+ \mu^- \gamma$	

- • • We do not use the following data for averages, fits, limits, etc. • • •
- <sup>1</sup> LEES 12E reports  $[\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(J/\psi(1S) \rightarrow \mu^+ \mu^-)] = (49.9 \pm 1.3 \pm 1.0) \times 10^{-3}$  keV which we divide by our best value  $B(J/\psi(1S) \rightarrow \mu^+ \mu^-) = (5.961 \pm 0.033) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> The value of  $\Gamma(e^+ e^-)$  quoted in BAI 98E is derived using  $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (32.4 \pm 2.6) \times 10^{-2}$  and  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1203 \pm 0.0038$ . Recalculated by us using  $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .
- <sup>3</sup> AUBERT 07AU reports  $[\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(J/\psi(1S) \rightarrow \pi^+ \pi^- \pi^0)] = 0.0186 \pm 0.0012 \pm 0.0011$  keV which we divide by our best value  $B(J/\psi(1S) \rightarrow \pi^+ \pi^- \pi^0) = (2.11 \pm 0.07) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>4</sup> AUBERT 05D reports  $[\Gamma(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(J/\psi(1S) \rightarrow \mu^+ \mu^-)] = 0.0450 \pm 0.0018 \pm 0.0022$  keV which we divide by our best value  $B(J/\psi(1S) \rightarrow \mu^+ \mu^-) = (5.961 \pm 0.033) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Superseded by LEES 12E.

WEIGHTED AVERAGE  
0.837 ± 0.025 (Error scaled by 1.3)



### $\Gamma(J/\psi(1S) \pi^0 \pi^0) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12} \Gamma_6 / \Gamma$
<b>0.425 ± 0.009 OUR FIT</b>					
<b>0.411 ± 0.008 ± 0.018</b>	3.6k ± 96	ADAM	06	CLEO $3.773 e^+ e^- \rightarrow \gamma \psi(2S)$	

### $\Gamma(J/\psi(1S) \eta) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{13} \Gamma_6 / \Gamma$
<b>78.6 ± 1.6 OUR FIT</b>					
<b>87 ± 9 OUR AVERAGE</b>					
83 ± 25 ± 5	14	<sup>1</sup> AUBERT	07AU	BABR $10.6 e^+ e^- \rightarrow J/\psi \pi^+ \pi^- \pi^0 \gamma$	
88 ± 6 ± 7	291 ± 24	ADAM	06	CLEO $3.773 e^+ e^- \rightarrow \gamma \psi(2S)$	

<sup>1</sup> AUBERT 07AU quotes  $\Gamma_{ee}^{\psi(2S)} \cdot B(\psi(2S) \rightarrow J/\psi \eta) \cdot B(J/\psi \rightarrow \mu^+ \mu^-) \cdot B(\eta \rightarrow \pi^+ \pi^- \pi^0) = 1.11 \pm 0.33 \pm 0.07$  eV.

### $\Gamma(J/\psi(1S) \pi^0) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{14} \Gamma_6 / \Gamma$
<b>&lt; 8</b>	90	< 37	ADAM	06	CLEO $3.773 e^+ e^- \rightarrow \gamma \psi(2S)$	

### $\Gamma(p \bar{p}) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{19} \Gamma_6 / \Gamma$
<b>0.674 ± 0.023 OUR FIT</b>					
<b>0.64 ± 0.04 OUR AVERAGE</b>					
0.67 ± 0.12 ± 0.02	43	LEES	13o	BABR $e^+ e^- \rightarrow p \bar{p} \gamma$	
0.74 ± 0.07 ± 0.04	142	LEES	13y	BABR $e^+ e^- \rightarrow p \bar{p} \gamma$	
0.579 ± 0.038 ± 0.036	2.7k	ANDREOTTI	07	E835 $p \bar{p} \rightarrow e^+ e^-$ , $J/\psi X$	
0.70 ± 0.17 ± 0.03	22	AUBERT	06b	$e^+ e^- \rightarrow p \bar{p} \gamma$	

### $\Gamma(\Lambda \bar{\Lambda}) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_{26} \Gamma_6 / \Gamma$
<b>1.5 ± 0.4 ± 0.1</b>	AUBERT	07BD	BABR $10.6 e^+ e^- \rightarrow \Lambda \bar{\Lambda} \gamma$	

### $\Gamma(2(\pi^+ \pi^- \pi^0)) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{59} \Gamma_6 / \Gamma$
<b>11.2 ± 3.3 ± 1.3</b>	43	AUBERT	06D	BABR $10.6 e^+ e^- \rightarrow 2(\pi^+ \pi^- \pi^0) \gamma$	

### $\Gamma(K^+ K^- 2(\pi^+ \pi^-)) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{73} \Gamma_6 / \Gamma$
<b>4.4 ± 2.1 ± 0.3</b>	26	AUBERT	06D	BABR $10.6 e^+ e^- \rightarrow K^+ K^- 2(\pi^+ \pi^-) \gamma$	

### $\Gamma(\pi^+ \pi^- K^+ K^-) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{68} \Gamma_6 / \Gamma$
<b>2.56 ± 0.42 ± 0.16</b>	85	AUBERT	07AK	BABR $10.6 e^+ e^- \rightarrow \pi^+ \pi^- K^+ K^- \gamma$	

### $\Gamma(\phi \phi(980) \rightarrow \pi^+ \pi^-) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{110} \Gamma_6 / \Gamma$
<b>0.347 ± 0.169 ± 0.003</b>	6 ± 3	<sup>1</sup> AUBERT	07AK	BABR $10.6 e^+ e^- \rightarrow \pi^+ \pi^- K^+ K^- \gamma$	

<sup>1</sup> AUBERT 07AK reports  $[\Gamma(\psi(2S) \rightarrow \phi \phi(980) \rightarrow \pi^+ \pi^-) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(\phi(1020) \rightarrow K^+ K^-)] = 0.17 \pm 0.08 \pm 0.02$  eV which we divide by our best value  $B(\phi(1020) \rightarrow K^+ K^-) = (48.9 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(\phi \pi^+ \pi^-) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{109} \Gamma_6 / \Gamma$
<b>0.57 ± 0.23 ± 0.01</b>	10	<sup>1</sup> AUBERT, BE	06D	BABR $10.6 e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$	

<sup>1</sup> AUBERT, BE 06D reports  $[\Gamma(\psi(2S) \rightarrow \phi \pi^+ \pi^-) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(\phi(1020) \rightarrow K^+ K^-)] = 0.28 \pm 0.11 \pm 0.02$  eV which we divide by our best value  $B(\phi(1020) \rightarrow K^+ K^-) = (48.9 \pm 0.5) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(2(\pi^+ \pi^-) \pi^0) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{17} \Gamma_6 / \Gamma$
<b>29.7 ± 2.2 ± 1.8</b>	410	AUBERT	07AU	BABR $10.6 e^+ e^- \rightarrow 2(\pi^+ \pi^-) \pi^0 \gamma$	

### $\Gamma(\omega \pi^+ \pi^-) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{64} \Gamma_6 / \Gamma$
<b>3.01 ± 0.84 ± 0.02</b>	37	<sup>1</sup> AUBERT	07AU	BABR $10.6 e^+ e^- \rightarrow \omega \pi^+ \pi^- \gamma$	

<sup>1</sup> AUBERT 07AU reports  $[\Gamma(\psi(2S) \rightarrow \omega \pi^+ \pi^-) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(\omega(782) \rightarrow \pi^+ \pi^- \pi^0)] = 2.69 \pm 0.73 \pm 0.16$  eV which we divide by our best value  $B(\omega(782) \rightarrow \pi^+ \pi^- \pi^0) = (89.2 \pm 0.7) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(2(\pi^+ \pi^-) \eta) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{62} \Gamma_6 / \Gamma$
<b>2.87 ± 1.41 ± 0.01</b>	16	<sup>1</sup> AUBERT	07AU	BABR $10.6 e^+ e^- \rightarrow 2(\pi^+ \pi^-) \eta \gamma$	

<sup>1</sup> AUBERT 07AU reports  $[\Gamma(\psi(2S) \rightarrow 2(\pi^+ \pi^-) \eta) \times \Gamma(\psi(2S) \rightarrow e^+ e^-) / \Gamma(\text{total})] \times [B(\eta \rightarrow 2\gamma)] = 1.13 \pm 0.55 \pm 0.08$  eV which we divide by our best value  $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) \times \Gamma(e^+ e^-) / \Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{80} \Gamma_6 / \Gamma$
<b>4.4 ± 1.3 ± 0.3</b>	32	AUBERT	07AU	BABR $10.6 e^+ e^- \rightarrow K^+ K^- \pi^+ \pi^- \pi^0 \gamma$	

See key on page 601

## Meson Particle Listings

 $\psi(2S)$ 

$\Gamma(K^+K^-\pi^+\pi^-\eta) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_{71}\Gamma_6/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3.04 ± 1.79 ± 0.02</b>	7	<sup>1</sup> AUBERT 07AU BABR	10.6	$e^+e^- \rightarrow K^+K^-\pi^+\pi^-\eta\gamma$	

<sup>1</sup> AUBERT 07AU reports  $[\Gamma(\psi(2S) \rightarrow K^+K^-\pi^+\pi^-\eta) \times \Gamma(\psi(2S) \rightarrow e^+e^-)/\Gamma_{\text{total}}] \times [B(\eta \rightarrow 2\gamma)] = 1.2 \pm 0.7 \pm 0.1$  eV which we divide by our best value  $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(K^+K^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_{98}\Gamma_6/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.147 ± 0.035 ± 0.005	66	<sup>1</sup> LEES	15J	BABR $e^+e^- \rightarrow K^+K^-\gamma$
0.197 ± 0.035 ± 0.005	66	<sup>2</sup> LEES	15J	BABR $e^+e^- \rightarrow K^+K^-\gamma$
0.35 ± 0.14 ± 0.03	11	<sup>3</sup> LEES	13Q	BABR $e^+e^- \rightarrow K^+K^-\gamma$

<sup>1</sup>  $\sin\phi > 0$ .

<sup>2</sup>  $\sin\phi < 0$ .

<sup>3</sup> Interference with non-resonant  $K^+K^-$  production not taken into account.

 $\psi(2S)$  BRANCHING RATIOS

$\Gamma(\text{hadrons})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.9785 ± 0.0013 OUR AVERAGE</b>				

0.9779 ± 0.0015	<sup>1</sup> BAI	02B	BES2 $e^+e^-$
0.981 ± 0.003	<sup>1</sup> LUTH	75	MRK1 $e^+e^-$

<sup>1</sup> Includes cascade decay into  $J/\psi(1S)$ .

$\Gamma(\text{virtual } \gamma \rightarrow \text{hadrons})/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.0173 ± 0.0014 OUR AVERAGE</b>	Error includes scale factor of 1.5.			

0.0166 ± 0.0010	<sup>1,2</sup> SETH	04	RVUE $e^+e^-$
0.0199 ± 0.0019	<sup>1</sup> BAI	02B	BES2 $e^+e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.029 ± 0.004	<sup>1</sup> LUTH	75	MRK1 $e^+e^-$
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<sup>1</sup> Included in  $\Gamma(\text{hadrons})/\Gamma_{\text{total}}$ .

<sup>2</sup> Using  $B(\psi(2S) \rightarrow \ell^+\ell^-) = (0.73 \pm 0.04)\%$  from RPP-2002 and  $R = 2.28 \pm 0.04$  determined by a fit to data from BAI 00 and BAI 02c.

$\Gamma(ggg)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>10.58 ± 1.62</b>	2.9 M	<sup>1</sup> LIBBY	09	CLEO $\psi(2S) \rightarrow \text{hadrons}$	

<sup>1</sup> Calculated using  $\Gamma(\gamma gg)/\Gamma(ggg) = 0.097 \pm 0.026 \pm 0.016$  from LIBBY 09,  $B(\psi(2S) \rightarrow XJ/\psi)$  relative and absolute branching fractions from MENDEZ 08,  $B(\psi(2S) \rightarrow \gamma\eta_c)$  from MITCHELL 09, and  $B(\psi(2S) \rightarrow \text{virtual } \gamma \rightarrow \text{hadrons})$ ,  $B(\psi(2S) \rightarrow \gamma\chi_{cJ})$ , and  $B(\psi(2S) \rightarrow \ell^+\ell^-)$  from PDG 08. The statistical error is negligible and the systematic error is largely uncorrelated with that of  $\Gamma(\gamma gg)/\Gamma_{\text{total}}$  LIBBY 09 measurement.

$\Gamma(\gamma gg)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.025 ± 0.288</b>	200 k	<sup>1</sup> LIBBY	09	CLEO $\psi(2S) \rightarrow \gamma + \text{hadrons}$	

<sup>1</sup> Calculated using  $\Gamma(\gamma gg)/\Gamma(ggg) = 0.097 \pm 0.026 \pm 0.016$  from LIBBY 09. The statistical error is negligible and the systematic error is largely uncorrelated with that of  $\Gamma(ggg)/\Gamma_{\text{total}}$  LIBBY 09 measurement.

$\Gamma(\gamma gg)/\Gamma(ggg)$					$\Gamma_4/\Gamma_3$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>9.7 ± 2.6 ± 1.6</b>	2.9 M	LIBBY	09	CLEO $\psi(2S) \rightarrow (\gamma +) \text{hadrons}$	

$\Gamma(\text{light hadrons})/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.154 ± 0.015</b>	<sup>1</sup> MENDEZ	08	CLEO $e^+e^- \rightarrow \psi(2S)$		

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.169 ± 0.026	<sup>2</sup> ADAM	05A	CLEO $e^+e^- \rightarrow \psi(2S)$
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<sup>1</sup> Uses  $B(\psi(2S) \rightarrow J/\psi X)$  from MENDEZ 08 and other branching fractions from PDG 07.

<sup>2</sup> Uses  $B(J/\psi X)$  from ADAM 05A,  $B(\chi_{cJ}\gamma)$ ,  $B(\eta_c\gamma)$  from ATHAR 04 and  $B(\ell^+\ell^-)$  from PDG 04. Superseded by MENDEZ 08.

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>78.9 ± 1.7 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

88 ± 13	<sup>1</sup> FELDMAN	77	RVUE $e^+e^-$
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<sup>1</sup> From an overall fit assuming equal partial widths for  $e^+e^-$  and  $\mu^+\mu^-$ . For a measurement of the ratio see the entry  $\Gamma(\mu^+\mu^-)/\Gamma(e^+e^-)$  below. Includes LUTH 75, HILGER 75, BURMESTER 77.

$\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$		$\Gamma_7/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	
<b>79 ± 9 OUR FIT</b>		

$\Gamma(\mu^+\mu^-)/\Gamma(e^+e^-)$				$\Gamma_7/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>1.00 ± 0.11 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.89 ± 0.16	BOYARSKI	75c	MRK1 $e^+e^-$
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$\Gamma(\tau^+\tau^-)/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT	
<b>31 ± 4 OUR FIT</b>				
<b>30.8 ± 2.1 ± 3.8</b>	<sup>1</sup> ABLIKIM	06w	BES $e^+e^- \rightarrow \psi(2S)$	

<sup>1</sup> Computed using PDG 02 value of  $B(\psi(2S) \rightarrow \text{hadrons}) = 0.9810 \pm 0.0030$  to estimate the total number of  $\psi(2S)$  events.

DECAYS INTO  $J/\psi(1S)$  AND ANYTHING

$\Gamma(J/\psi(1S) \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.610 ± 0.006 OUR FIT</b>					
<b>0.55 ± 0.07 OUR AVERAGE</b>					

0.51 ± 0.12	BRANDELIK	79c	DASP $e^+e^- \rightarrow \mu^+\mu^-X$
0.57 ± 0.08	ABRAMS	75B	MRK1 $e^+e^- \rightarrow \mu^+\mu^-X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.6254 ± 0.0016 ± 0.0155	1.1M	<sup>1</sup> MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-X$
0.5950 ± 0.0015 ± 0.0190	151k	ADAM	05A	CLEO Repl. by MENDEZ 08

<sup>1</sup> Not independent from other measurements of MENDEZ 08.

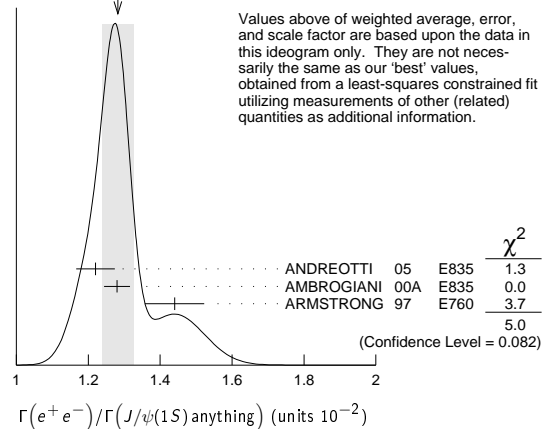
$\Gamma(e^+e^-)/\Gamma(J/\psi(1S) \text{ anything})$					$\Gamma_6/\Gamma_9 = \Gamma_6/(\Gamma_{11} + \Gamma_{12} + \Gamma_{13} + 0.339\Gamma_{133} + 0.192\Gamma_{134})$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.294 ± 0.026 OUR FIT</b>					

**1.28 ± 0.04 OUR AVERAGE** Error includes scale factor of 1.6. See the ideogram below.

1.22 ± 0.02 ± 0.05	5097 ± 73	<sup>1</sup> ANDREOTTI	05	E835 $p\bar{p} \rightarrow \psi(2S) \rightarrow e^+e^-$
1.28 ± 0.03 ± 0.02		<sup>1</sup> AMBROGIANI	00A	E835 $p\bar{p} \rightarrow \psi(2S)$
1.44 ± 0.08 ± 0.02		<sup>1</sup> ARMSTRONG	97	E760 $p\bar{p} \rightarrow \psi(2S)$

<sup>1</sup> Using  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0593 \pm 0.0010$ .

WEIGHTED AVERAGE  
1.28 ± 0.04 (Error scaled by 1.6)



$\Gamma(\mu^+\mu^-)/\Gamma(J/\psi(1S) \text{ anything})$					$\Gamma_7/\Gamma_9 = \Gamma_7/(\Gamma_{11} + \Gamma_{12} + \Gamma_{13} + 0.339\Gamma_{133} + 0.192\Gamma_{134})$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.0130 ± 0.0014 OUR FIT</b>					

0.014 ± 0.003	HILGER	75	SPEC $e^+e^-$
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$\Gamma(J/\psi(1S) \text{ neutrals})/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.2514 ± 0.0033 OUR FIT</b>				

$\Gamma(J/\psi(1S) \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.3449 ± 0.0030 OUR FIT</b>					

**0.348 ± 0.005 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

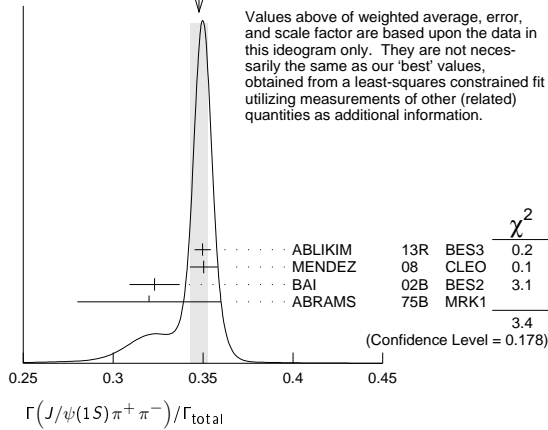
0.3498 ± 0.0002 ± 0.0045	20M	ABLIKIM	13R	BES3 $\psi(2S) \rightarrow J/\psi \pi^+\pi^-$
0.3504 ± 0.0007 ± 0.0077	565k	MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-\pi^+\pi^-$
0.323 ± 0.014		BAI	02B	BES2 $e^+e^-$
0.32 ± 0.04		ABRAMS	75B	MRK1 $e^+e^- \rightarrow J/\psi \pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.3354 ± 0.0014 ± 0.0110	60k	<sup>1</sup> ADAM	05A	CLEO Repl. by MENDEZ 08
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<sup>1</sup> Not independent from other values reported by ADAM 05A.

## Meson Particle Listings

 $\psi(2S)$ WEIGHTED AVERAGE  
0.348±0.005 (Error scaled by 1.3) $\Gamma(e^+e^-)/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_6/\Gamma_{11}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0229±0.0005 OUR FIT</b>			
<b>0.0252±0.0028±0.0011</b>	<sup>1</sup> AUBERT	02B	BABR $e^+e^-$

<sup>1</sup> Using  $B(J/\psi(1S) \rightarrow e^+e^-) = 0.0593 \pm 0.0010$ .

 $\Gamma(\mu^+\mu^-)/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_7/\Gamma_{11}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.0229±0.0025 OUR FIT</b>			
<b>0.0224±0.0029 OUR AVERAGE</b>			
0.0216±0.0026±0.0014	<sup>1</sup> AUBERT	02B	BABR $e^+e^-$
0.0327±0.0077±0.0072	<sup>1</sup> GRIBUSHIN	96	FMPs 515 $\pi^-Be \rightarrow 2\mu X$

<sup>1</sup> Using  $B(J/\psi(1S) \rightarrow \mu^+\mu^-) = 0.0588 \pm 0.0010$ .

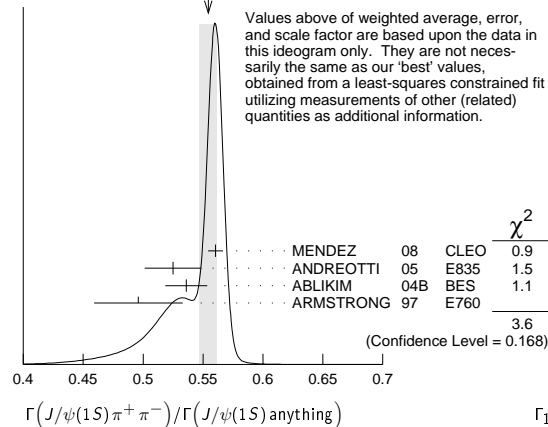
 $\Gamma(\tau^+\tau^-)/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_8/\Gamma_{11}$ 

VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT
<b>8.9 ±1.1 OUR FIT</b>			
<b>8.73±1.39±1.57</b>	BAI	02	BES $e^+e^-$

 $\Gamma(J/\psi(1S)\pi^+\pi^-)/\Gamma(J/\psi(1S)\text{anything})$  $\Gamma_{11}/\Gamma_9$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.5653±0.0026 OUR FIT</b>				
<b>0.554 ±0.008 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
0.5604±0.0009±0.0062	565k	MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-\pi^+\pi^-$
0.525 ±0.009 ±0.022	4k	ANDREOTTI	05	E835 $\psi(2S) \rightarrow J/\psi X$
0.536 ±0.007 ±0.016	20k	<sup>1,2</sup> ABLIKIM	04B	BES $\psi(2S) \rightarrow J/\psi X$
0.496 ±0.037		ARMSTRONG	97	E760 $\bar{p}p \rightarrow \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.5637±0.0027±0.0046	60k	ADAM	05A	CLEO Repl. by MENDEZ 08

<sup>1</sup> From a fit to the  $J/\psi$  recoil mass spectra.  
<sup>2</sup> ABLIKIM 04B quotes  $B(\psi(2S) \rightarrow J/\psi X) / B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-)$ .

WEIGHTED AVERAGE  
0.554±0.008 (Error scaled by 1.3) $\Gamma(J/\psi(1S)\text{neutrals})/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_{10}/\Gamma_{11} = (0.9761\Gamma_{12} + 0.719\Gamma_{13} + 0.339\Gamma_{133} + 0.192\Gamma_{134})/\Gamma_{11}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.729±0.008 OUR FIT</b>			
<b>0.73 ±0.09</b>	TANENBAUM	76	MRK1 $e^+e^-$

 $\Gamma(J/\psi(1S)\pi^0\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{12}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1816±0.0031 OUR FIT</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.1769±0.0008±0.0053	61k	<sup>1</sup> MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-2\pi^0$
0.1652±0.0014±0.0058	13.4k	<sup>2</sup> ADAM	05A	CLEO Repl. by MENDEZ 08

<sup>1</sup> Not independent from other measurements of MENDEZ 08.  
<sup>2</sup> Not independent from other values reported by ADAM 05A.

 $\Gamma(J/\psi(1S)\pi^0\pi^0)/\Gamma(J/\psi(1S)\text{anything})$  $\Gamma_{12}/\Gamma_9$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.2977±0.0031 OUR FIT</b>				
<b>0.320 ±0.012 OUR AVERAGE</b>				
0.300 ±0.008 ±0.022	1655 ± 44	ANDREOTTI	05	E835 $\psi(2S) \rightarrow J/\psi X$
0.328 ±0.013 ±0.008		AMBROGIANI	00A	E835 $p\bar{p} \rightarrow \psi(2S)$
0.323 ±0.033		ARMSTRONG	97	E760 $\bar{p}p \rightarrow \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.2829±0.0012±0.0056	61k	MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-2\pi^0$
0.2776±0.0025±0.0043	13.4k	ADAM	05A	CLEO Repl. by MENDEZ 08

 $\Gamma(J/\psi(1S)\pi^0\pi^0)/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_{12}/\Gamma_{11}$ 

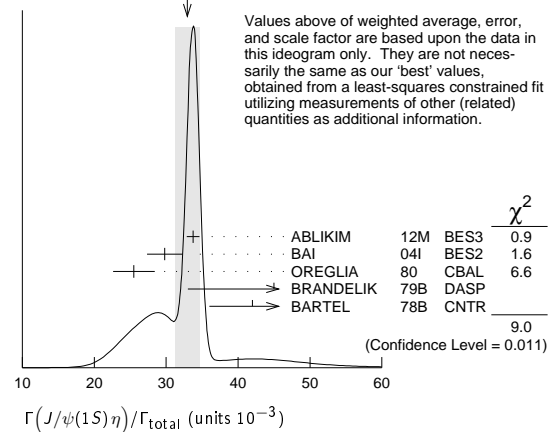
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.527 ±0.008 OUR FIT</b>				
<b>0.513 ±0.022 OUR AVERAGE</b>				Error includes scale factor of 2.2.
0.5047±0.0022±0.0102	61k	MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-2\pi^0$
0.570 ±0.009 ±0.026	14k	<sup>1</sup> ABLIKIM	04B	BES $\psi(2S) \rightarrow J/\psi X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.4924±0.0047±0.0086	73k	<sup>2,3</sup> ADAM	05A	CLEO Repl. by MENDEZ 08
0.571 ±0.018 ±0.044		<sup>4</sup> ANDREOTTI	05	E835 $\psi(2S) \rightarrow J/\psi X$
0.53 ±0.06		TANENBAUM	76	MRK1 $e^+e^-$
0.64 ±0.15		<sup>5</sup> HILGER	75	SPEC $e^+e^-$

<sup>1</sup> From a fit to the  $J/\psi$  recoil mass spectra.  
<sup>2</sup> Not independent from other values reported by ADAM 05A.  
<sup>3</sup> Using 13,217  $J/\psi\pi^0\pi^0$  and 60,010  $J/\psi\pi^+\pi^-$  events.  
<sup>4</sup> Not independent from other values reported by ANDREOTTI 05.  
<sup>5</sup> Ignoring the  $J/\psi(1S)\eta$  and  $J/\psi(1S)\gamma\gamma$  decays.

 $\Gamma(J/\psi(1S)\eta)/\Gamma_{\text{total}}$  $\Gamma_{13}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>33.6 ± 0.5 OUR FIT</b>				
<b>32.9 ± 1.7 OUR AVERAGE</b>				Error includes scale factor of 2.1. See the ideogram below.
33.75 ± 0.17 ± 0.86	68.2k	ABLIKIM	12M	BES3 $e^+e^- \rightarrow \ell^+\ell^-2\gamma$
29.8 ± 0.9 ± 2.3	5.7k	BAI	04I	BES2 $J/\psi\gamma\gamma$
25.5 ± 2.9	386	<sup>1</sup> OREGLIA	80	CBAL $e^+e^- \rightarrow J/\psi 2\gamma$
45 ± 12	17	<sup>2</sup> BRANDELIK	79B	DASP $e^+e^- \rightarrow J/\psi 2\gamma$
42 ± 6	164	<sup>2</sup> BARTEL	78B	CNTR $e^+e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
34.3 ± 0.4 ± 0.9	18.4k	<sup>3</sup> MENDEZ	08	CLEO $\psi(2S) \rightarrow \ell^+\ell^-\eta$
32.5 ± 0.6 ± 1.1	2.8k	<sup>4</sup> ADAM	05A	CLEO Repl. by MENDEZ 08
43 ± 8	44	TANENBAUM	76	MRK1 $e^+e^-$

<sup>1</sup> Recalculated by us using  $B(J/\psi(1S) \rightarrow \ell^+\ell^-) = 0.1181 \pm 0.0020$ .  
<sup>2</sup> Recalculated by us using  $B(J/\psi(1S) \rightarrow \mu^+\mu^-) = 0.0588 \pm 0.0010$ .  
<sup>3</sup> Not independent from other measurements of MENDEZ 08.  
<sup>4</sup> Not independent from other values reported by ADAM 05A.

WEIGHTED AVERAGE  
32.9±1.7 (Error scaled by 2.1)

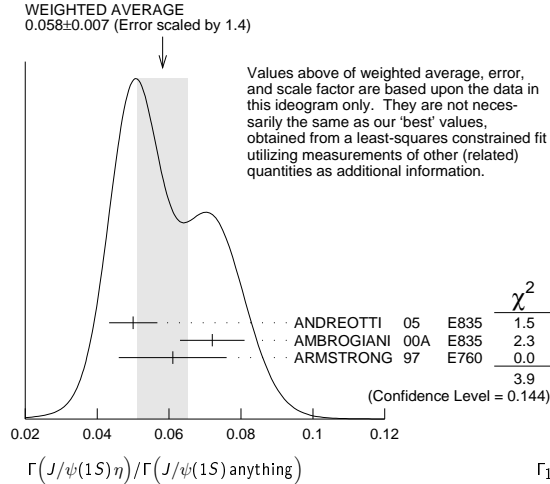


See key on page 601

## Meson Particle Listings

 $\psi(2S)$  $\Gamma(J/\psi(1S)\eta)/\Gamma(J/\psi(1S)\text{anything})$  $\Gamma_{13}/\Gamma_9$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0551±0.0008 OUR FIT</b>				
<b>0.058 ±0.007 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
0.050 ±0.006 ±0.003	298 ± 20	ANDREOTTI 05	E835	$\psi(2S) \rightarrow J/\psi X$
0.072 ±0.009		AMBROGIANI 00A	E835	$p\bar{p} \rightarrow \psi(2S)$
0.061 ±0.015		ARMSTRONG 97	E760	$\bar{p}p \rightarrow \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0549±0.0006±0.0009	18.4k	<sup>1</sup> MENDEZ 08	CLEO	$\psi(2S) \rightarrow \ell^+ \ell^- \eta$
0.0546±0.0010±0.0007	2.8k	ADAM 05A	CLEO	Repl. by MENDEZ 08
<sup>1</sup> Not independent from other measurements of MENDEZ 08.				

 $\Gamma(J/\psi(1S)\eta)/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_{13}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0974±0.0014 OUR FIT</b>				
<b>0.0979±0.0018 OUR AVERAGE</b>				
0.0979±0.0010±0.0015	18.4k	MENDEZ 08	CLEO	$\psi(2S) \rightarrow \ell^+ \ell^- \eta$
0.098 ±0.005 ±0.010	2k	<sup>1</sup> ABLIKIM 04b	BES	$\psi(2S) \rightarrow J/\psi X$
0.091 ±0.021		<sup>2</sup> HIMEL 80	MRK2	$e^+e^- \rightarrow \psi(2S)X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0968±0.0019±0.0013	2.8k	<sup>3</sup> ADAM 05A	CLEO	Repl. by MENDEZ 08
0.095 ±0.007 ±0.007		<sup>4</sup> ANDREOTTI 05	E835	$\psi(2S) \rightarrow J/\psi X$
<sup>1</sup> From a fit to the $J/\psi$ recoil mass spectra.				
<sup>2</sup> The value for $B(\psi(2S) \rightarrow J/\psi(1S)\eta)$ reported in HIMEL 80 is derived using $B(\psi(2S)) \rightarrow J/\psi(1S)\pi^+\pi^- = (33 \pm 3)\%$ and $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.138 \pm 0.018$ . Calculated by us using $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = (0.1181 \pm 0.0020)$ .				
<sup>3</sup> Not independent from other values reported by ADAM 05A.				
<sup>4</sup> Not independent from other values reported by ANDREOTTI 05.				

 $\Gamma(J/\psi(1S)\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{14}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>12.68±0.32 OUR AVERAGE</b>				
12.6 ±0.2 ±0.3	4.1k	ABLIKIM 12M	BES3	$e^+e^- \rightarrow \ell^+ \ell^- 2\gamma$
13.3 ±0.8 ±0.3	530	MENDEZ 08	CLEO	$\psi(2S) \rightarrow \ell^+ \ell^- 2\gamma$
14.3 ±1.4 ±1.2	280	BAI 04i	BES2	$\psi(2S) \rightarrow J/\psi \gamma \gamma$
14 ±6	7	HIMEL 80	MRK2	$e^+e^-$
9 ±2 ±1	23	<sup>1</sup> OREGLIA 80	CBAL	$\psi(2S) \rightarrow J/\psi 2\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13 ±1 ±1	88	ADAM 05A	CLEO	Repl. by MENDEZ 08
<sup>1</sup> Recalculated by us using $B(J/\psi(1S) \rightarrow \ell^+ \ell^-) = 0.1181 \pm 0.0020$ .				

 $\Gamma(J/\psi(1S)\pi^0)/\Gamma(J/\psi(1S)\text{anything})$  $\Gamma_{14}/\Gamma_9 = \Gamma_{14}/(\Gamma_{11} + \Gamma_{12} + \Gamma_{13} + 0.339\Gamma_{133} + 0.192\Gamma_{134})$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.213±0.012±0.003	527	<sup>1</sup> MENDEZ 08	CLEO	$e^+e^- \rightarrow J/\psi \gamma \gamma$
0.22 ±0.02 ±0.01		<sup>2</sup> ADAM 05A	CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow J/\psi \gamma \gamma$
<sup>1</sup> Not independent from other values reported by MENDEZ 08. Supersedes ADAM 05A.				
<sup>2</sup> Not independent from other values reported by ADAM 05A.				

 $\Gamma(J/\psi(1S)\pi^0)/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_{14}/\Gamma_{11}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.380±0.022±0.005	527	<sup>1</sup> MENDEZ 08	CLEO	$e^+e^- \rightarrow J/\psi \gamma \gamma$
0.39 ±0.04 ±0.01		<sup>2</sup> ADAM 05A	CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow J/\psi \gamma \gamma$
<sup>1</sup> Not independent from other values reported by MENDEZ 08. Supersedes ADAM 05A.				
<sup>2</sup> Not independent from other values reported by ADAM 05A.				

## HADRONIC DECAYS

 $\Gamma(\pi^0 h_c(1P))/\Gamma_{\text{total}}$  $\Gamma_{15}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.6±1.3 OUR AVERAGE</b>				
9.0±1.5±1.3	3k	<sup>1</sup> GE 11	CLEO	$\psi(2S) \rightarrow \pi^0 \text{ anything}$
8.4±1.3±1.0	11k	ABLIKIM 10b	BES3	$\psi(2S) \rightarrow \pi^0 h_c$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	92 <sup>+23</sup> <sub>-22</sub>	ADAMS 09	CLEO	$\psi(2S) \rightarrow 2\pi^+ 2\pi^- 2\pi^0$
seen	1282	DOBBS 08A	CLEO	$\psi(2S) \rightarrow \pi^0 \eta_c \gamma$
seen	168 ± 40	ROSNER 05	CLEO	$\psi(2S) \rightarrow \pi^0 \eta_c \gamma$

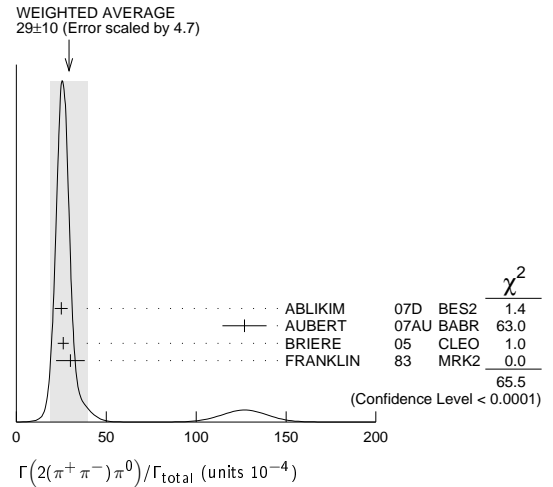
<sup>1</sup> Assuming a width  $\Gamma(h_c(1P)) = 0.86 \text{ MeV} \equiv \Gamma_0$ , a measured dependence of the central value of  $B = (7.6 + 1.4 \times \Gamma(h_c(1P))/\Gamma_0) \times 10^{-4}$ , and with a systematic error that accounts for the width variation range 0.43–1.29 MeV.

 $\Gamma(3(\pi^+\pi^-)\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{16}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>35±16</b>	6	FRANKLIN 83	MRK2	$e^+e^- \rightarrow \text{hadrons}$

 $\Gamma(2(\pi^+\pi^-)\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{17}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>29 ±10 OUR AVERAGE</b>				Error includes scale factor of 4.7. See the ideogram below.
24.9± 0.7±3.6	2173	ABLIKIM 07D	BES2	$e^+e^- \rightarrow \psi(2S)$
127 ±12 ±2	410	<sup>1</sup> AUBERT 07Au	BABR	$10.6 e^+e^- \rightarrow 2(\pi^+\pi^-)\pi^0 \gamma$
26.1± 0.7±3.0	1703	BRIERE 05	CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow 2(\pi^+\pi^-)\pi^0$
30 ± 8	42	FRANKLIN 83	MRK2	$e^+e^-$
<sup>1</sup> AUBERT 07Au reports $[\Gamma(\psi(2S) \rightarrow 2(\pi^+\pi^-)\pi^0)/\Gamma_{\text{total}}] \times [\Gamma(\psi(2S) \rightarrow e^+e^-)] = (297 \pm 22 \pm 18) \times 10^{-4} \text{ keV}$ which we divide by our best value $\Gamma(\psi(2S) \rightarrow e^+e^-) = 2.34 \pm 0.04 \text{ keV}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

 $\Gamma(\rho a_2(1320))/\Gamma_{\text{total}}$  $\Gamma_{18}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.55±0.73±0.47</b>					
3.36±0.09±0.25	112 ± 31		BAI 04c	BES2	$\psi(2S) \rightarrow 2(\pi^+\pi^-)\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<2.3	90		BAI 98J	BES	$e^+e^-$

 $\Gamma(\rho\bar{\rho})/\Gamma_{\text{total}}$  $\Gamma_{19}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.88±0.09 OUR FIT</b>				
<b>3.00±0.13 OUR AVERAGE</b>				Error includes scale factor of 1.1.
3.08±0.05±0.18	4.5k	<sup>1</sup> DOBBS 14		$e^+e^- \rightarrow \psi(2S) \rightarrow p\bar{p}$
3.36±0.09±0.25	1.6k	ABLIKIM 07c	BES	$e^+e^- \rightarrow \psi(2S) \rightarrow p\bar{p}$
2.87±0.12±0.15	557	PEDLAR 05	CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow p\bar{p}$
1.4 ±0.8	4	BRANDELIK 79c	DASP	$e^+e^- \rightarrow \psi(2S) \rightarrow p\bar{p}$
2.3 ±0.7		FELDMAN 77	MRK1	$e^+e^- \rightarrow \psi(2S) \rightarrow p\bar{p}$
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.				

 $\Gamma(\rho\bar{\rho})/\Gamma(J/\psi(1S)\pi^+\pi^-)$  $\Gamma_{19}/\Gamma_{11}$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>8.35±0.28 OUR FIT</b>			
<b>6.98±0.49±0.97</b>	BAI 01	BES	$e^+e^- \rightarrow \psi(2S) \rightarrow p\bar{p}$

 $\Gamma(\Delta^{++}\bar{\Delta}^{--})/\Gamma_{\text{total}}$  $\Gamma_{20}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>12.8±1.0±3.4</b>	157	<sup>1</sup> BAI 01	BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
<sup>1</sup> Estimated using $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = 0.310 \pm 0.028$ .				

# Meson Particle Listings

## $\psi(2S)$

$\Gamma(\Lambda\overline{\Lambda}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 0.29	90	<sup>1</sup> ABLIKIM	13F BES3	$\psi(2S) \rightarrow p\overline{p}\pi^+\pi^-\gamma\gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<12	90	<sup>2</sup> ABLIKIM	07H BES2	$e^+e^- \rightarrow \psi(2S)$	

<sup>1</sup> Using  $B(\Lambda \rightarrow \pi^-p) = 63.9\%$  and  $B(\pi^0 \rightarrow \gamma\gamma) = 98.8\%$ .  
<sup>2</sup> Using  $B(\Lambda \rightarrow \pi^-p) = 63.9\%$  and  $B(\eta \rightarrow \gamma\gamma) = 39.4\%$ .

$\Gamma(\Lambda\overline{\Lambda}\eta)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>2.48±0.34±0.19</b>	60		<sup>1</sup> ABLIKIM	13F BES3	$\psi(2S) \rightarrow p\overline{p}\pi^+\pi^-\gamma\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<4.9	90		<sup>2</sup> ABLIKIM	07H BES2	$e^+e^- \rightarrow \psi(2S)$

<sup>1</sup> Using  $B(\Lambda \rightarrow \pi^-p) = 63.9\%$  and  $B(\eta \rightarrow \gamma\gamma) = 39.31\%$ .  
<sup>2</sup> Using  $B(\Lambda \rightarrow \pi^-p) = 63.9\%$ .

$\Gamma(\Lambda\overline{p}K^+)/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>1.0±0.1±0.1</b>	74.0	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow p\overline{p}K^+\pi^-$	

$\Gamma(\Lambda\overline{p}K^+\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>1.8±0.3±0.3</b>	45.8	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow p\overline{p}K^+\pi^+\pi^-\pi^-$	

$\Gamma(\Lambda\overline{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{25}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>2.8±0.4±0.5</b>	73.4	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow p\overline{p}2(\pi^+\pi^-)$	

$\Gamma(\Lambda\overline{\Lambda})/\Gamma_{\text{total}}$					$\Gamma_{26}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>3.57±0.18 OUR AVERAGE</b>					
3.75±0.09±0.23	1.9k		<sup>1</sup> DOBBS	14	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
3.39±0.20±0.32	337		ABLIKIM	07c BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
6.4±1.8±0.1			<sup>2</sup> AUBERT	07BD BABR	10.6 $e^+e^- \rightarrow \Lambda\overline{\Lambda}\gamma$
3.28±0.23±0.25	208		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.81±0.20±0.27	80		<sup>3</sup> BAI	01 BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
< 4	90		FELDMAN	77 MRK1	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.  
<sup>2</sup> AUBERT 07BD reports  $[\Gamma(\psi(2S) \rightarrow \Lambda\overline{\Lambda})/\Gamma_{\text{total}}] \times [\Gamma(\psi(2S) \rightarrow e^+e^-)] = (15 \pm 4 \pm 1) \times 10^{-4}$  keV which we divide by our best value  $\Gamma(\psi(2S) \rightarrow e^+e^-) = 2.34 \pm 0.04$  keV. Our first error is their experiment's error and our second error is the systematic error from using our best value.  
<sup>3</sup> Estimated using  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.310 \pm 0.028$ .

$\Gamma(\Lambda\Sigma^+\pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>1.40±0.03±0.13</b>	2.8k	ABLIKIM	13W BES3	$\psi(2S) \rightarrow \text{hadrons}$	

$\Gamma(\Lambda\Sigma^-\pi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>1.54±0.04±0.13</b>	2.8k	ABLIKIM	13W BES3	$\psi(2S) \rightarrow \text{hadrons}$	

$\Gamma(\Sigma^0\overline{p}K^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-5}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>1.67±0.13±0.12</b>	276	<sup>1</sup> ABLIKIM	13D BES3	$\psi(2S) \rightarrow \gamma\Lambda\overline{p}K^+$	

<sup>1</sup> Using  $B(\Lambda \rightarrow p\pi^-) = 63.9\%$ , and  $B(\Sigma^0 \rightarrow \Lambda\gamma) = 100\%$ .

$\Gamma(\Sigma^+\Sigma^-)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>2.51±0.21 OUR AVERAGE</b>					
2.51±0.15±0.16	281		<sup>1</sup> DOBBS	14	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
2.57±0.44±0.68	35		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.

$\Gamma(\Sigma^0\Sigma^0)/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>2.32±0.16 OUR AVERAGE</b>					
2.25±0.11±0.16	439		<sup>1</sup> DOBBS	14	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
2.35±0.36±0.32	59		ABLIKIM	07c BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
2.63±0.35±0.21	58		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.2±0.4±0.4	8		<sup>2</sup> BAI	01 BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.  
<sup>2</sup> Estimated using  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.310 \pm 0.028$ .

$\Gamma(\Sigma(1385)^+\overline{\Sigma}(1385)^-)/\Gamma_{\text{total}}$					$\Gamma_{32}/\Gamma$
VALUE (units $10^{-5}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>11±3±3</b>	14	<sup>1</sup> BAI	01 BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$	

<sup>1</sup> Estimated using  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.310 \pm 0.028$ .

$\Gamma(\Xi^-\Xi^+)/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>2.64±0.18 OUR AVERAGE</b>					
2.66±0.12±0.20	548		<sup>1</sup> DOBBS	14	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
3.03±0.40±0.32	67		ABLIKIM	07c BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
2.38±0.30±0.21	63		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.94±0.27±0.15	12		<sup>2</sup> BAI	01 BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
<2	90		FELDMAN	77 MRK1	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.  
<sup>2</sup> Estimated using  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.310 \pm 0.028$ .

$\Gamma(\Xi^0\Xi^0)/\Gamma_{\text{total}}$					$\Gamma_{34}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>2.07±0.23 OUR AVERAGE</b>					
2.02±0.19±0.15	112		<sup>1</sup> DOBBS	14	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
2.75±0.64±0.61	19		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.

$\Gamma(\Xi(1530)^0\Xi(1530)^0)/\Gamma_{\text{total}}$					$\Gamma_{35}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>5.2±0.3±3.2</b> <b>-1.2</b>		527	<sup>1</sup> ABLIKIM	13s BES3	$\psi(2S) \rightarrow \eta p\overline{p}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<32	90		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
< 8.1	90		<sup>2</sup> BAI	01 BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> With  $N(1535)$  decaying to  $p\eta$ .  
<sup>2</sup> Estimated using  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.310 \pm 0.028$ .

$\Gamma(K^-\Lambda\Xi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{36}/\Gamma$
VALUE (units $10^{-5}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>3.86±0.27±0.32</b>	236	ABLIKIM	15i BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow K^-\Lambda\Xi^+ + \text{c.c.}$	

$\Gamma(\Xi(1690)^-\Xi^+ \rightarrow K^-\Lambda\Xi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{37}/\Gamma$
VALUE (units $10^{-6}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>5.21±1.48±0.57</b>	74	ABLIKIM	15i BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow K^-\Lambda\Xi^+ + \text{c.c.}$	

$\Gamma(\Xi(1820)^-\Xi^+ \rightarrow K^-\Lambda\Xi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{38}/\Gamma$
VALUE (units $10^{-6}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>12.03±2.94±1.22</b>	136	ABLIKIM	15i BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow K^-\Lambda\Xi^+ + \text{c.c.}$	

$\Gamma(K^-\Sigma^0\Xi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{39}/\Gamma$
VALUE (units $10^{-5}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>3.67±0.33±0.28</b>	142	ABLIKIM	15i BES3	$e^+e^- \rightarrow \psi(2S) \rightarrow K^-\Sigma^0\Xi^+ + \text{c.c.}$	

$\Gamma(\Omega^-\overline{\Omega}^+)/\Gamma_{\text{total}}$					$\Gamma_{40}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.47±0.09±0.05</b>		27	<sup>1</sup> DOBBS	14	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.5	90		ABLIKIM	12q BES2	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
<1.6	90		PEDLAR	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$
<0.73	90		<sup>2</sup> BAI	01 BES	$e^+e^- \rightarrow \psi(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.  
<sup>2</sup> Estimated using  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.310 \pm 0.028$ .

$\Gamma(\pi^0\rho\overline{\rho})/\Gamma_{\text{total}}$					$\Gamma_{41}/\Gamma$
VALUE (units $10^{-4}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>1.53±0.07 OUR AVERAGE</b>					
1.65±0.03±0.15	4.5k		ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\overline{p}\pi^0$
1.54±0.06±0.06	948		ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \pi^0 p\overline{p}$
1.32±0.10±0.15	256		<sup>1</sup> ABLIKIM	05E BES2	$e^+e^- \rightarrow \psi(2S) \rightarrow p\overline{p}\gamma\gamma$
1.4±0.5	9		FRANKLIN	83 MRK2	$e^+e^-$

<sup>1</sup> Computed using  $B(\pi^0 \rightarrow \gamma\gamma) = (98.80 \pm 0.03)\%$ .

$\Gamma(N(940)\overline{p} + \text{c.c.} \rightarrow \pi^0\rho\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{42}/\Gamma$
VALUE (units $10^{-5}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b>6.42±0.20±1.78</b> <b>-1.28</b>	1.9k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\overline{p}\pi^0$	

<sup>1</sup> From a fit of  $\pi^0\rho\overline{p}$  data to eight distinct intermediate  $N\overline{p}$  resonant states.

See key on page 601

## Meson Particle Listings

 $\psi(2S)$  $\Gamma(N(1440)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{43}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>7.3^{+1.7}_{-1.5}</math> OUR AVERAGE</b>				Error includes scale factor of 2.5.
$3.58 \pm 0.25^{+1.59}_{-0.84}$	1.1k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
$8.1 \pm 0.7 \pm 0.3$	474	<sup>2</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \pi^0 p\bar{p}$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				
<sup>2</sup> From a fit of the $p\bar{p}$ and $p\pi^0$ mass distributions to a combination of $N(1440)\bar{p}$ , $\pi^0 f_0(2100)$ , and two other broad, unestablished resonances.				

 $\Gamma(N(1520)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{44}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.64 \pm 0.05^{+0.22}_{-0.17}</math></b>	0.2k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				

 $\Gamma(N(1535)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{45}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.47 \pm 0.28^{+0.99}_{-0.97}</math></b>	0.7k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				

 $\Gamma(N(1650)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{46}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.76 \pm 0.28^{+1.37}_{-1.66}</math></b>	1.1k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				

 $\Gamma(N(1720)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{47}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.79 \pm 0.10^{+0.24}_{-0.71}</math></b>	0.5k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				

 $\Gamma(N(2300)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{48}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.62 \pm 0.28^{+1.12}_{-0.64}</math></b>	0.9k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				

 $\Gamma(N(2570)\bar{p} + \text{c.c.} \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{49}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.13 \pm 0.08^{+0.40}_{-0.30}</math></b>	0.8k	<sup>1</sup> ABLIKIM	13A BES3	$\psi(2S) \rightarrow p\bar{p}\pi^0$
<sup>1</sup> From a fit of $\pi^0 p\bar{p}$ data to eight distinct intermediate $N\bar{p}$ resonant states.				

 $\Gamma(\pi^0 f_0(2100) \rightarrow \pi^0 p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{50}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 0.4 \pm 0.1</math></b>	76	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \pi^0 p\bar{p}$
<sup>1</sup> From a fit of the $p\bar{p}$ and $p\pi^0$ mass distributions to a combination of $N_1^*(1440)\bar{p}$ , $\pi^0 f_0(2100)$ , and two other broad, unestablished resonances.				

 $\Gamma(\eta p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{51}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.0 \pm 0.4</math> OUR AVERAGE</b>				Error includes scale factor of 2.1. See the ideogram below.
$6.4 \pm 0.2 \pm 0.6$	679	<sup>1</sup> ABLIKIM	13s BES3	$\psi(2S) \rightarrow \eta p\bar{p}$
$5.6 \pm 0.6 \pm 0.3$	154	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \eta p\bar{p}$
$5.8 \pm 1.1 \pm 0.7$	44.8 $\pm$ 8.5	<sup>2</sup> ABLIKIM	05e BES2	$e^+ e^- \rightarrow \psi(2S) \rightarrow p\bar{p}\gamma\gamma$
$8 \pm 3 \pm 3$	9.8	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow p\bar{p}\pi^+ \pi^- \pi^0$

<sup>1</sup> With  $N(1535)$  decaying to  $p\eta$ .<sup>2</sup> Computed using  $B(\eta \rightarrow \gamma\gamma) = (39.43 \pm 0.26)\%$ . $\Gamma(\eta f_0(2100) \rightarrow \eta p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{52}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.2 \pm 0.4 \pm 0.1</math></b>	31	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \eta p\bar{p}$
<sup>1</sup> From a fit of the $p\bar{p}$ and $p\eta$ distributions to a combination of $N^*(1535)\bar{p}$ and $\eta f_0(2100)$ .				

 $\Gamma(N(1535)\bar{p} \rightarrow \eta p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{53}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.4 \pm 0.6 \pm 0.3</math></b>	123	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \eta p\bar{p}$
<sup>1</sup> From a fit of the $p\bar{p}$ and $p\eta$ distributions to a combination of $N^*(1535)\bar{p}$ and $\eta f_0(2100)$ .				

 $\Gamma(\omega p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{54}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.69 \pm 0.21</math> OUR AVERAGE</b>				
$0.6 \pm 0.2 \pm 0.2$	21.2	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow p\bar{p}\pi^+ \pi^- \pi^0$
$0.8 \pm 0.3 \pm 0.1$	14.9 $\pm$ 0.1	<sup>1</sup> BAI	03B BES	$\psi(2S) \rightarrow p\bar{p}\pi^+ \pi^- \pi^0$
<sup>1</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = 0.305 \pm 0.016$ .				

 $\Gamma(\phi p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{55}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.24</math></b>	90	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow p\bar{p}K^+ K^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<0.26$	90	<sup>1</sup> BAI	03B BES	$\psi(2S) \rightarrow K^+ K^- p\bar{p}$
<sup>1</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = 0.305 \pm 0.016$ .				

 $\Gamma(\pi^+ \pi^- p\bar{p})/\Gamma_{\text{total}}$   $\Gamma_{56}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.0 \pm 0.4</math> OUR AVERAGE</b>				
$5.9 \pm 0.2 \pm 0.4$	904.5	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow p\bar{p}\pi^+ \pi^-$
$8 \pm 2$		<sup>1</sup> TANENBAUM	78 MRK1	$e^+ e^-$
<sup>1</sup> Assuming entirely strong decay.				

 $\Gamma(p\bar{p}\pi^- \text{ or c.c.})/\Gamma_{\text{total}}$   $\Gamma_{57}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.48 \pm 0.17</math> OUR AVERAGE</b>				
$2.45 \pm 0.11 \pm 0.21$	851	ABLIKIM	06i BES2	$e^+ e^- \rightarrow p\pi^- X$
$2.52 \pm 0.12 \pm 0.22$	849	ABLIKIM	06i BES2	$e^+ e^- \rightarrow \bar{p}\pi^+ X$

 $\Gamma(p\bar{p}\pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{58}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.18 \pm 0.50 \pm 0.50</math></b>	135 $\pm$ 21	ABLIKIM	06i BES2	$e^+ e^- \rightarrow p\pi^- \pi^0 X$

 $\Gamma(\eta\pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{60}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.6</math></b>	90	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow 2(\pi^+ \pi^-) \pi^0$

 $\Gamma(\eta\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{61}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>9.5 \pm 0.7 \pm 1.5</math></b>		<sup>1</sup> BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow \text{had}^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$10.3 \pm 0.8 \pm 1.4$	201.7	<sup>2</sup> BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow \eta 3\pi(\eta \rightarrow \gamma\gamma)$
$8.1 \pm 1.4 \pm 1.6$	50.0	<sup>2</sup> BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow \eta 3\pi(\eta \rightarrow 3\pi)$

<sup>1</sup> Average of  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow 3\pi$ .<sup>2</sup> Not independent from other values reported by BRIERE 05. $\Gamma(2(\pi^+ \pi^-)\eta)/\Gamma_{\text{total}}$   $\Gamma_{62}/\Gamma$ 

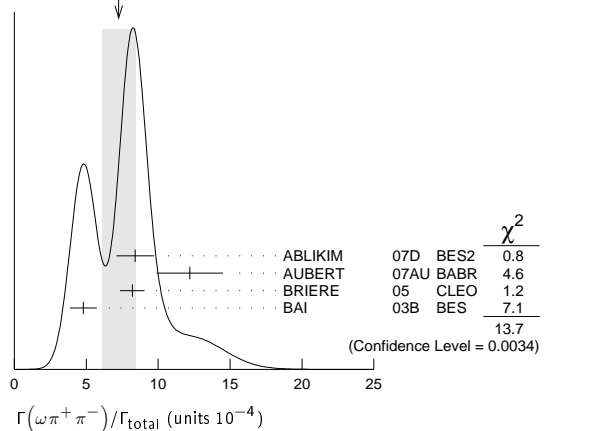
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.2 \pm 0.6 \pm 0.1</math></b>	16	<sup>1</sup> AUBERT	07AU BABR	$10.6 e^+ e^- \rightarrow 2(\pi^+ \pi^-)\eta\gamma$
<sup>1</sup> AUBERT 07AU quotes $\Gamma_{ee}^{\psi(2S)} \cdot B(\psi(2S) \rightarrow 2(\pi^+ \pi^-)\eta) \cdot B(\eta \rightarrow \gamma\gamma) = 1.2 \pm 0.7 \pm 0.1 \text{ eV}$ .				

 $\Gamma(\eta' \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{63}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.5 \pm 1.6 \pm 1.3</math></b>	12.8	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow \text{had}^+$

 $\Gamma(\omega\pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{64}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>7.3 \pm 1.2</math> OUR AVERAGE</b>				Error includes scale factor of 2.1. See the ideogram below.
$8.4 \pm 0.5 \pm 1.2$	386	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$
$12.2 \pm 2.2 \pm 0.7$	37	<sup>1</sup> AUBERT	07AU BABR	$10.6 e^+ e^- \rightarrow \omega\pi^+ \pi^- \gamma$
$8.2 \pm 0.5 \pm 0.7$	391	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow 2(\pi^+ \pi^-) \pi^0$
$4.8 \pm 0.6 \pm 0.7$	100 $\pm$ 22	<sup>2</sup> BAI	03B BES	$\psi(2S) \rightarrow 2(\pi^+ \pi^-) \pi^0$

WEIGHTED AVERAGE  
7.3 $\pm$ 1.2 (Error scaled by 2.1)



$\Gamma(\omega K_S^0(1430)^+ K^- + \text{c.c.})/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.1 ± 1.2 OUR AVERAGE</b>				
6.39 ± 1.50 ± 0.78	128	ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K_S^0 K^- \pi^+$
5.86 ± 1.61 ± 0.83	143	ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K^+ K^- \pi^0$

 $\Gamma(\omega \bar{K}^*(892)^0 K^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>16.8 ± 2.5 ± 1.6</b>	356	ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K_S^0 K^- \pi^+$

 $\Gamma(\omega \bar{K}_S^0(1430)^0 K^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.82 ± 2.08 ± 0.72</b>	116	ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K_S^0 K^- \pi^+$

 $\Gamma(\omega X(1440) \rightarrow \omega K_S^0 K^- \pi^+ + \text{c.c.})/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.60 ± 0.27 ± 0.24</b>	109	<sup>1</sup> ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K_S^0 K^- \pi^+$

<sup>1</sup> X(1440) compatible with  $\eta(1405)$  and  $\eta(1475)$ . A  $f_1(1420)$  is also possible.

 $\Gamma(\omega X(1440) \rightarrow \omega K^+ K^- \pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.09 ± 0.20 ± 0.16</b>	82	<sup>1</sup> ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K^+ K^- \pi^0$

<sup>1</sup> X(1440) compatible with  $\eta(1405)$  and  $\eta(1475)$ . A  $f_1(1420)$  is also possible.

 $\Gamma(\omega f_1(1285) \rightarrow \omega K_S^0 K^- \pi^+ + \text{c.c.})/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.302 ± 0.098 ± 0.027</b>	22	<sup>1</sup> ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K_S^0 K^- \pi^+$

<sup>1</sup> Statistical significance 4.5  $\sigma$ . This measurement is equivalent to a limit of  $< 0.478 \times 10^{-5}$  at 90% C.L.

 $\Gamma(\omega f_1(1285) \rightarrow \omega K^+ K^- \pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.125 ± 0.070 ± 0.013</b>	10	<sup>1</sup> ABLIKIM	13M BES3	$\psi(2S) \rightarrow \omega K^+ K^- \pi^0$

<sup>1</sup> Statistical significance 3.2  $\sigma$ . This measurement is equivalent to a limit of  $< 0.221 \times 10^{-5}$  at 90% C.L.

 $\Gamma(3(\pi^+ \pi^-))/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.5 ± 2.0 OUR AVERAGE</b>				Error includes scale factor of 2.8.
5.45 ± 0.42 ± 0.87	671	ABLIKIM	05H BES2	$e^+ e^- \rightarrow \psi(2S) \rightarrow 3(\pi^+ \pi^-)$

1.5 ± 1.0 <sup>1</sup> TANENBAUM 78 MRK1  $e^+ e^-$

<sup>1</sup> Assuming entirely strong decay.

 $\Gamma(p\bar{p}\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.3 ± 0.4 ± 0.6</b>	434.9	BRIERE	05 CLEO	$e^+ e^- \rightarrow \psi(2S) \rightarrow p\bar{p}\pi^+\pi^-\pi^0$

 $\Gamma(K^+ K^-)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.48 ± 0.23 ± 0.39</b>	1.3k		<sup>1</sup> METREVELI	12	$\psi(2S) \rightarrow K^+ K^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

6.2 ± 1.5 ± 0.2	66	<sup>2,3</sup> LEES	15J BABR	$e^+ e^- \rightarrow K^+ K^- \gamma$
8.3 ± 1.5 ± 0.2	66	<sup>3,4</sup> LEES	15J BABR	$e^+ e^- \rightarrow K^+ K^- \gamma$
6.3 ± 0.6 ± 0.3		<sup>5</sup> DOBBS	06A CLEO	$e^+ e^-$
10 ± 7		<sup>5</sup> BRANDELIK	79c DASP	$e^+ e^-$
< 5	90	FELDMAN	77 MRK1	$e^+ e^-$

<sup>1</sup> Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration.

<sup>2</sup>  $\sin\phi > 0$ .

<sup>3</sup> Using  $\Gamma(\psi(2S) \rightarrow e^+ e^-) = (2.37 \pm 0.04) \text{ keV}$ .

<sup>4</sup>  $\sin\phi < 0$ .

<sup>5</sup> Interference with non-resonant  $K^+ K^-$  production not taken into account.

 $\Gamma(K_S^0 K_L^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.34 ± 0.33 OUR AVERAGE</b>				
5.28 ± 0.25 ± 0.34	478 ± 23	<sup>1</sup> METREVELI	12	$\psi(2S) \rightarrow K_S^0 K_L^0$
5.8 ± 0.8 ± 0.4		DOBBS	06A CLEO	$e^+ e^-$
5.24 ± 0.47 ± 0.48	156 ± 14	<sup>2</sup> BAI	04B BES2	$\psi(2S) \rightarrow K_S^0 K_L^0 \rightarrow \pi^+ \pi^- X$

<sup>1</sup> Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration.

<sup>2</sup> Using  $B(K_S^0 \rightarrow \pi^+ \pi^-) = 0.6860 \pm 0.0027$ .

 $\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$ 

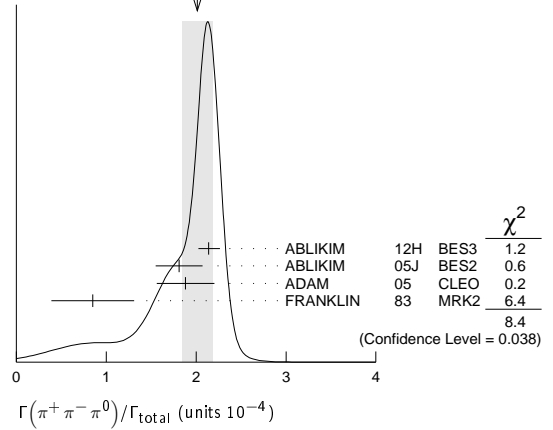
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.01 ± 0.17 OUR AVERAGE</b>				Error includes scale factor of 1.7. See the ideogram below.

2.14 ± 0.03 ± 0.12	7k	<sup>1</sup> ABLIKIM	12H BES3	$e^+ e^- \rightarrow \psi(2S)$
1.81 ± 0.18 ± 0.19	260 ± 19	<sup>2</sup> ABLIKIM	05J BES2	$e^+ e^- \rightarrow \psi(2S)$
1.88 ± 0.16 ± 0.28	194	ADAM	05 CLEO	$e^+ e^- \rightarrow \psi(2S)$
0.85 ± 0.46	4	FRANKLIN	83 MRK2	$e^+ e^- \rightarrow \text{hadrons}$

<sup>1</sup> From  $\psi(2S) \rightarrow \pi^+ \pi^- \pi^0$  events directly. The quoted systematic error includes a contribution of 4% (added in quadrature) from the uncertainty on the number of  $\psi(2S)$  events.

<sup>2</sup> From a PW analysis of  $\psi(2S) \rightarrow \pi^+ \pi^- \pi^0$ .

WEIGHTED AVERAGE  
2.01 ± 0.17 (Error scaled by 1.7)

 $\Gamma(\rho(2150)\pi \rightarrow \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.94 ± 0.25 ± 1.15</b>	<sup>1</sup> ABLIKIM	05J BES2	$\psi(2S) \rightarrow \rho(2150)\pi \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> From a PW analysis of  $\psi(2S) \rightarrow \pi^+ \pi^- \pi^0$ .

 $\Gamma(\rho(770)\pi \rightarrow \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.32 ± 0.12 OUR AVERAGE</b>					Error includes scale factor of 1.8.
0.51 ± 0.07 ± 0.11			<sup>1</sup> ABLIKIM	05J BES2	$\psi(2S) \rightarrow \rho(770)\pi \rightarrow \pi^+ \pi^- \pi^0$

0.24 ± 0.08 ± 0.02 22 ADAM 05 CLEO  $e^+ e^- \rightarrow \psi(2S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.83	90	1	FRANKLIN	83 MRK2	$e^+ e^-$
< 10	90		BARTEL	76 CNTR	$e^+ e^-$
< 10	90		<sup>2</sup> ABRAMS	75 MRK1	$e^+ e^-$

<sup>1</sup> From a PW analysis of  $\psi(2S) \rightarrow \pi^+ \pi^- \pi^0$ .

<sup>2</sup> Final state  $\rho^0 \pi^0$ .

 $\Gamma(\pi^+ \pi^-)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.78 ± 0.26 OUR AVERAGE</b>					

0.76 ± 0.25 ± 0.06	30	<sup>1</sup> METREVELI	12	$\psi(2S) \rightarrow \pi^+ \pi^-$
8 ± 5		BRANDELIK	79c DASP	$e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 2.1	90		DOBBS	06A CLEO	$e^+ e^- \rightarrow \psi(2S)$
< 5	90		FELDMAN	77 MRK1	$e^+ e^-$

<sup>1</sup> Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration. Using  $\psi(3770) \rightarrow \pi^+ \pi^-$  for continuum subtraction.

 $\Gamma(K_1(1400)^\pm K^\mp)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.1</b>	90	<sup>1</sup> BAI	99c BES	$e^+ e^-$

<sup>1</sup> Assuming  $B(K_1(1400) \rightarrow K^* \pi) = 0.94 \pm 0.06$

 $\Gamma(K_S^0(1430)^\pm K^\mp)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.12 ± 0.62 ± 1.13</b>	251 ± 22	ABLIKIM	12L BES3	$e^+ e^- \rightarrow \psi(2S)$

 $\Gamma(K^+ K^- \pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.07 ± 0.16 ± 0.26</b>	0.9k		ABLIKIM	12L BES3	$e^+ e^- \rightarrow \psi(2S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 8.9	90	1	FRANKLIN	83 MRK2	$e^+ e^- \rightarrow \text{hadrons}$
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# Meson Particle Listings

## $\psi(2S)$

$\Gamma(K^+K^*(892)^- + c.c.)/\Gamma_{\text{total}}$					$\Gamma_{107}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.9 \pm 0.4</math></b>	<b>OUR AVERAGE</b> Error includes scale factor of 1.2.				
$3.18 \pm 0.30^{+0.26}_{-0.31}$		0.2k	ABLIKIM	12L BES3	$e^+e^- \rightarrow \psi(2S)$
$2.9^{+1.3}_{-1.7} \pm 0.4$		$9.6 \pm 4.2$	ABLIKIM	05i BES2	$e^+e^- \rightarrow \psi(2S)$
$1.3^{+1.0}_{-0.7} \pm 0.3$		7	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<5.4		90	FRANKLIN	83 MRK2	$e^+e^- \rightarrow \text{hadrons}$

$\Gamma(K^*(892)^0\bar{K}^0 + c.c.)/\Gamma_{\text{total}}$					$\Gamma_{108}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>10.9 \pm 2.0</math></b>	<b>OUR AVERAGE</b>				
$13.3^{+2.4}_{-2.8} \pm 1.7$		$65.6 \pm 9.0$	ABLIKIM	05i BES2	$e^+e^- \rightarrow \psi(2S)$
$9.2^{+2.7}_{-2.2} \pm 0.9$		25	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(K^+K^*(892)^- + c.c.)/\Gamma(K^*(892)^0\bar{K}^0 + c.c.)$					$\Gamma_{107}/\Gamma_{108}$
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.16 \pm 0.06</math></b>	<b>OUR AVERAGE</b>				
$0.22^{+0.10}_{-0.14}$			ABLIKIM	05i BES2	$e^+e^- \rightarrow \psi(2S)$
$0.14^{+0.08}_{-0.06}$			ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\phi\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{109}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.17 \pm 0.29</math></b>	<b>OUR AVERAGE</b> Error includes scale factor of 1.7.				
$2.44 \pm 0.96 \pm 0.04$		$10 \pm 4$	<sup>1,2</sup> AUBERT	07Ak BABR	$10.6 e^+e^- \rightarrow \pi^+\pi^-K^+K^-\gamma$
$0.9 \pm 0.2 \pm 0.1$		47.6	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow K^+K^-\pi^+\pi^-$
$1.5 \pm 0.2 \pm 0.2$		$51.5 \pm 8.3$	<sup>3</sup> BAI	03B BES	$\psi(2S) \rightarrow K^+K^-\pi^+\pi^-$
<sup>1</sup> AUBERT 07Ak reports $[\Gamma(\psi(2S) \rightarrow \phi\pi^+\pi^-)/\Gamma_{\text{total}}] \times [\Gamma(\psi(2S) \rightarrow e^+e^-)] = (0.57 \pm 0.22 \pm 0.04) \times 10^{-3}$ keV which we divide by our best value $\Gamma(\psi(2S) \rightarrow e^+e^-) = 2.34 \pm 0.04$ keV. Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using $B(\phi \rightarrow K^+K^-) = (49.3 \pm 0.6)\%$ .					
<sup>3</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.305 \pm 0.016$ .					

$\Gamma(\phi\phi_0(980) \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{110}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.60 \pm 0.25</math></b>	<b>OUR AVERAGE</b> Error includes scale factor of 1.2.				
$1.45 \pm 0.70 \pm 0.02$		$6 \pm 3$	<sup>1,2</sup> AUBERT	07Ak BABR	$10.6 e^+e^- \rightarrow \pi^+\pi^-K^+K^-\gamma$
$0.6 \pm 0.2 \pm 0.1$		$18.4 \pm 6.4$	<sup>3</sup> BAI	03B BES	$\psi(2S) \rightarrow K^+K^-\pi^+\pi^-$
<sup>1</sup> AUBERT 07Ak reports $[\Gamma(\psi(2S) \rightarrow \phi\phi_0(980) \rightarrow \pi^+\pi^-)/\Gamma_{\text{total}}] \times [\Gamma(\psi(2S) \rightarrow e^+e^-)] = (0.34 \pm 0.16 \pm 0.04) \times 10^{-3}$ keV which we divide by our best value $\Gamma(\psi(2S) \rightarrow e^+e^-) = 2.34 \pm 0.04$ keV. Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using $B(\phi \rightarrow K^+K^-) = (49.3 \pm 0.6)\%$ .					
<sup>3</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.305 \pm 0.016$ .					

$\Gamma(2(K^+K^-))/\Gamma_{\text{total}}$					$\Gamma_{111}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.6 \pm 0.1 \pm 0.1</math></b>		59.2	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow 2(K^+K^-)$

$\Gamma(\phi K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{112}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.70 \pm 0.16</math></b>	<b>OUR AVERAGE</b>				
$0.8 \pm 0.2 \pm 0.1$		36.8	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow 2(K^+K^-)$
$0.6 \pm 0.2 \pm 0.1$		$16.1 \pm 5.0$	<sup>1</sup> BAI	03B BES	$\psi(2S) \rightarrow 2(K^+K^-)$
<sup>1</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.305 \pm 0.016$ .					

$\Gamma(2(K^+K^-)\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{113}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 0.2 \pm 0.2</math></b>		44.7	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow 2(K^+K^-)\pi^0$

$\Gamma(\phi\eta)/\Gamma_{\text{total}}$					$\Gamma_{114}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.10 \pm 0.31</math></b>	<b>OUR AVERAGE</b>				
$3.14 \pm 0.23 \pm 0.23$		0.2k	ABLIKIM	12L BES3	$e^+e^- \rightarrow \psi(2S)$
$2.0^{+1.5}_{-1.1} \pm 0.4$		6	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$
$3.3 \pm 1.1 \pm 0.5$		17	ABLIKIM	04K BES	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\phi\eta')/\Gamma_{\text{total}}$					$\Gamma_{115}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.1 \pm 1.4 \pm 0.7</math></b>		8	<sup>1</sup> ABLIKIM	04K BES	$e^+e^- \rightarrow \psi(2S)$
<sup>1</sup> Calculated combining $\eta' \rightarrow \gamma\rho$ and $\eta\pi^+\pi^-$ channels.					

$\Gamma(\omega\eta')/\Gamma_{\text{total}}$					$\Gamma_{116}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.2 \pm 2.4 \pm 0.7</math></b>		4	<sup>1</sup> ABLIKIM	04K BES	$e^+e^- \rightarrow \psi(2S)$
<sup>1</sup> Calculated combining $\eta' \rightarrow \gamma\rho$ and $\eta\pi^+\pi^-$ channels.					

$\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{117}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.1 \pm 0.6</math></b>	<b>OUR AVERAGE</b>				
$2.5^{+1.2}_{-1.0} \pm 0.2$		14	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$
$1.87^{+0.68}_{-0.62} \pm 0.28$		14	ABLIKIM	04L BES	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\rho\eta')/\Gamma_{\text{total}}$					$\Gamma_{118}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.87 \pm 1.64 \pm 0.33</math></b>		2	ABLIKIM	04L BES	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\rho\eta)/\Gamma_{\text{total}}$					$\Gamma_{119}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.2 \pm 0.6</math></b>	<b>OUR AVERAGE</b> Error includes scale factor of 1.1.				
$3.0^{+1.1}_{-0.9} \pm 0.2$		18	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$
$1.78^{+0.67}_{-0.62} \pm 0.17$		13	ABLIKIM	04L BES	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\omega\eta)/\Gamma_{\text{total}}$					$\Gamma_{120}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1</b>		90	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<3.1		90	ABLIKIM	04K BES	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\phi\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{121}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.04</b>		90	ABLIKIM	12L BES3	$e^+e^- \rightarrow \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.7		90	ADAM	05 CLEO	$e^+e^- \rightarrow \psi(2S)$
<0.4		90	ABLIKIM	04K BES	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\eta_c\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{122}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0</b>		90	PEDLAR	07 CLEO	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\rho\bar{\rho}K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{123}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.7 \pm 0.6 \pm 0.4</math></b>		30.1	BRIERE	05 CLEO	$e^+e^- \rightarrow \psi(2S) \rightarrow \rho\bar{\rho}K^+K^-$

$\Gamma(\bar{\Lambda}nK_S^0 + c.c.)/\Gamma_{\text{total}}$					$\Gamma_{124}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.81 \pm 0.11 \pm 0.14</math></b>		50	<sup>1</sup> ABLIKIM	08c BES2	$e^+e^- \rightarrow J/\psi$
<sup>1</sup> Using $B(\bar{\Lambda} \rightarrow \bar{p}\pi^+) = 63.9\%$ and $B(K_S^0 \rightarrow \pi^+\pi^-) = 69.2\%$ .					

$\Gamma(\phi f_2'(1525))/\Gamma_{\text{total}}$					$\Gamma_{125}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.44 \pm 0.12 \pm 0.11</math></b>		$20 \pm 6$	BAI	04c	$\psi(2S) \rightarrow 2(K^+K^-)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.45		90	BAI	98J BES	$e^+e^- \rightarrow 2(K^+K^-)$

$\Gamma(\Theta(1540)\bar{\Theta}(1540) \rightarrow K_S^0\rho K^-\bar{\pi} + c.c.)/\Gamma_{\text{total}}$					$\Gamma_{126}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.88</b>		90	BAI	04G BES2	$e^+e^-$

$\Gamma(\Theta(1540)K^-\bar{\pi} \rightarrow K_S^0\rho K^-\bar{\pi})/\Gamma_{\text{total}}$					$\Gamma_{127}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.0</b>		90	BAI	04G BES2	$e^+e^-$

$\Gamma(\Theta(1540)K_S^0\bar{\rho} \rightarrow K_S^0\bar{\rho}K^+n)/\Gamma_{\text{total}}$					$\Gamma_{128}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.70</b>		90	BAI	04G BES2	$e^+e^-$

$\Gamma(\bar{\Theta}(1540)K^+n \rightarrow K_S^0\bar{\rho}K^+n)/\Gamma_{\text{total}}$					$\Gamma_{129}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.6</b>		90	BAI	04G BES2	$e^+e^-$

$\Gamma(\bar{\Theta}(1540)K_S^0\rho \rightarrow K_S^0\rho K^-\bar{\pi})/\Gamma_{\text{total}}$					$\Gamma_{130}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.60</b>		90	BAI	04G BES2	$e^+e^-$

See key on page 601

## Meson Particle Listings

 $\psi(2S)$ 

$\Gamma(K_S^0 K_S^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{131}/\Gamma$
VALUE (units $10^{-4}$ )				
<b>&lt;0.046</b>	<sup>1</sup> BAI	04D	BES $e^+e^-$	

<sup>1</sup> Forbidden by CP.

## RADIATIVE DECAYS

$\Gamma(\gamma\chi_{c0}(1P))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{132}/\Gamma$
VALUE (units $10^{-2}$ )				
<b>9.99±0.27 OUR FIT</b>				
<b>9.2 ±0.4 OUR AVERAGE</b>				
9.22±0.11±0.46	72600	ATHAR	04 CLEO $e^+e^- \rightarrow \gamma X$	
9.9 ±0.5 ±0.8		<sup>1</sup> GAISER	86 CBAL $e^+e^- \rightarrow \gamma X$	
7.2 ±2.3		<sup>1</sup> BIDDICK	77 CNTR $e^+e^- \rightarrow \gamma X$	
7.5 ±2.6		<sup>1</sup> WHITAKER	76 MRK1 $e^+e^-$	

<sup>1</sup> Angular distribution  $(1+\cos^2\theta)$  assumed.

$\Gamma(\gamma\chi_{c1}(1P))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{133}/\Gamma$
VALUE (units $10^{-2}$ )				
<b>9.55±0.31 OUR FIT</b>				
<b>8.9 ±0.5 OUR AVERAGE</b>				
9.07±0.11±0.54	76700	ATHAR	04 CLEO $e^+e^- \rightarrow \gamma X$	
9.0 ±0.5 ±0.7		<sup>1</sup> GAISER	86 CBAL $e^+e^- \rightarrow \gamma X$	
7.1 ±1.9		<sup>2</sup> BIDDICK	77 CNTR $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> Angular distribution  $(1-0.189\cos^2\theta)$  assumed.<sup>2</sup> Valid for isotropic distribution of the photon.

$\Gamma(\gamma\chi_{c2}(1P))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{134}/\Gamma$
VALUE (units $10^{-2}$ )				
<b>9.11±0.31 OUR FIT</b>				
<b>8.8 ±0.5 OUR AVERAGE</b>				
9.33±0.14±0.61	79300	ATHAR	04 CLEO $e^+e^- \rightarrow \gamma X$	
8.0 ±0.5 ±0.7		<sup>1</sup> GAISER	86 CBAL $e^+e^- \rightarrow \gamma X$	
7.0 ±2.0		<sup>2</sup> BIDDICK	77 CNTR $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> Angular distribution  $(1-0.052\cos^2\theta)$  assumed.<sup>2</sup> Valid for isotropic distribution of the photon.

$[\Gamma(\gamma\chi_{c0}(1P)) + \Gamma(\gamma\chi_{c1}(1P)) + \Gamma(\gamma\chi_{c2}(1P))]/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_{132}+\Gamma_{133}+\Gamma_{134})/\Gamma$
VALUE				

• • • We do not use the following data for averages, fits, limits, etc. • • •

27.6±0.3±2.0 <sup>1</sup> ATHAR 04 CLEO  $e^+e^- \rightarrow \gamma X$ <sup>1</sup> Not independent from ATHAR 04 measurements of  $B(\gamma\chi_{cJ})$ .

$\Gamma(\gamma\chi_{c0}(1P))/\Gamma(\gamma\chi_{c1}(1P))$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{132}/\Gamma_{133}$
VALUE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.02±0.01±0.07	<sup>1</sup> ATHAR	04	CLEO $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> Not independent from ATHAR 04 measurements of  $B(\gamma\chi_{cJ})$ .

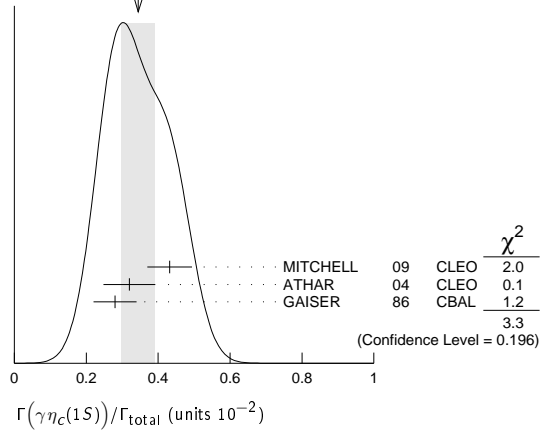
$\Gamma(\gamma\chi_{c2}(1P))/\Gamma(\gamma\chi_{c1}(1P))$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{134}/\Gamma_{133}$
VALUE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.03±0.02±0.03	<sup>1</sup> ATHAR	04	CLEO $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> Not independent from ATHAR 04 measurements of  $B(\gamma\chi_{cJ})$ .

$\Gamma(\gamma\chi_{c0}(1P))/\Gamma(\gamma\chi_{c2}(1P))$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{132}/\Gamma_{134}$
VALUE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.99±0.02±0.08	<sup>1</sup> ATHAR	04	CLEO $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> Not independent from ATHAR 04 measurements of  $B(\gamma\chi_{cJ})$ .

$\Gamma(\gamma\eta_c(1S))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{135}/\Gamma$
VALUE (units $10^{-2}$ )				
<b>0.34 ±0.05 OUR AVERAGE</b>				
0.432±0.016±0.060		MITCHELL	09 CLEO $e^+e^- \rightarrow \gamma X$	
0.32 ±0.04 ±0.06	2560	<sup>1</sup> ATHAR	04 CLEO $e^+e^- \rightarrow \gamma X$	
0.28 ±0.06		<sup>2</sup> GAISER	86 CBAL $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> ATHAR 04 used  $\Gamma_{\eta_c(1S)} = 24.8 \pm 4.9$  MeV to obtain this result.<sup>2</sup> GAISER 86 used  $\Gamma_{\eta_c(1S)} = 11.5 \pm 4.5$  MeV to obtain this result.WEIGHTED AVERAGE  
0.34±0.05 (Error scaled by 1.3)

$\Gamma(\gamma\eta_c(2S))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{136}/\Gamma$
VALUE (units $10^{-4}$ )				
<b>7±2±4</b>	<sup>1</sup> ABLIKIM	12G	BES3 $\psi(2S) \rightarrow \gamma K^0 K \pi, K K \pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 8	90	<sup>2</sup> CRONIN-HEN..10	CLEO $\psi(2S) \rightarrow \gamma K \bar{K} \pi$	
<20	90	ATHAR	04 CLEO $e^+e^- \rightarrow \gamma X$	
20–130	95	EDWARDS	82c CBAL $e^+e^- \rightarrow \gamma X$	

<sup>1</sup> ABLIKIM 12G reports  $[\Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}] \times [B(\eta_c(2S) \rightarrow K\bar{K}\pi)] = (1.30 \pm 0.20 \pm 0.30) \times 10^{-9}$  which we divide by our best value  $B(\eta_c(2S) \rightarrow K\bar{K}\pi) = (1.9 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>2</sup> CRONIN-HENNESSY 10 reports  $[\Gamma(\psi(2S) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}] \times [B(\eta_c(2S) \rightarrow K\bar{K}\pi)] < 14.5 \times 10^{-6}$  which we divide by our best value  $B(\eta_c(2S) \rightarrow K\bar{K}\pi) = 1.9 \times 10^{-2}$ . This measurement assumes  $\Gamma(\eta_c(2S)) = 14$  MeV. CRONIN-HENNESSY 10 gives the analytic dependence of limits on width.

$\Gamma(\gamma\pi^0)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{137}/\Gamma$
VALUE (units $10^{-5}$ )				
<b>1.58±0.40±0.13</b>	37	ABLIKIM	10F BES3 $\psi(2S) \rightarrow \gamma\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 5	90	PEDLAR	09 CLE3 $\psi(2S) \rightarrow \gamma X$	
<5400	95	<sup>1</sup> LIBERMAN	75 SPEC $e^+e^-$	
< $1 \times 10^4$	90	WIJK	75 DASP $e^+e^-$	

<sup>1</sup> Restated by us using  $B(\psi(2S) \rightarrow \mu^+\mu^-) = 0.0077$ .

$\Gamma(\gamma\eta'(958))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{138}/\Gamma$
VALUE (units $10^{-4}$ )				
<b>1.23±0.06 OUR AVERAGE</b>				
1.26±0.03±0.08	2226	<sup>1</sup> ABLIKIM	10F BES3 $\psi(2S) \rightarrow 3\gamma\pi^+\pi^-, 2\gamma\pi^+\pi^-$	
1.19±0.08±0.03		PEDLAR	09 CLE3 $\psi(2S) \rightarrow \gamma X$	
1.24±0.27±0.15	23	ABLIKIM	06R BES3 $e^+e^- \rightarrow \psi(2S)$	
1.54±0.31±0.20	~43	BAI	98F BES $\psi(2S) \rightarrow \pi^+\pi^-\pi^0, \pi^+\pi^-\pi^+\pi^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 60	90	<sup>2</sup> BRAUNSCH...	77 DASP $e^+e^-$	
< 11	90	<sup>3</sup> BARTEL	76 CNTR $e^+e^-$	

<sup>1</sup> Combining the results from  $\eta' \rightarrow \pi^+\pi^-\eta$  and  $\eta' \rightarrow \pi^+\pi^-\gamma$  decay modes.<sup>2</sup> Restated by us using total decay width 228 keV.<sup>3</sup> The value is normalized to the branching ratio for  $\Gamma(J/\psi(1S)\eta)/\Gamma_{\text{total}}$ .

$\Gamma(\gamma f_2(1270))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{139}/\Gamma$
VALUE (units $10^{-4}$ )				
<b>2.73±0.29 OUR AVERAGE</b>				
2.84±0.15±0.03	1.9k	<sup>1,2</sup> DOBBS	15 $\psi(2S) \rightarrow \gamma\pi\pi$	
2.12±0.19±0.32		<sup>3,4</sup> BAI	03c BES $\psi(2S) \rightarrow \gamma\pi\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.08±0.19±0.33	200.6 ± 18.8	<sup>3</sup> BAI	03c BES $\psi(2S) \rightarrow \gamma\pi^+\pi^-$	
2.90±1.08±1.07	29.9 ± 11.1	<sup>3</sup> BAI	03c BES $\psi(2S) \rightarrow \gamma\pi^0\pi^0$	

<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.<sup>2</sup> DOBBS 15 reports  $[\Gamma(\psi(2S) \rightarrow \gamma f_2(1270))/\Gamma_{\text{total}}] \times [B(f_2(1270) \rightarrow \pi\pi)] = (2.39 \pm 0.09 \pm 0.09) \times 10^{-4}$  which we divide by our best value  $B(f_2(1270) \rightarrow \pi\pi) = (84.2^{+2.9}_{-0.9}) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>3</sup> Normalized to  $B(\psi(2S) \rightarrow J/\psi\pi^+\pi^-) = 0.305 \pm 0.016$ .<sup>4</sup> Combining the results from  $\pi^+\pi^-$  and  $\pi^0\pi^0$  decay modes.

# Meson Particle Listings

## $\psi(2S)$

$\Gamma(\gamma f_0(1370) \rightarrow \gamma K \bar{K})/\Gamma_{\text{total}}$					$\Gamma_{140}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	COMMENT		
<b><math>3.1 \pm 1.0 \pm 1.4</math></b>	175	<sup>1</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K \bar{K}$	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					

$\Gamma(\gamma f_0(1500))/\Gamma_{\text{total}}$					$\Gamma_{141}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	COMMENT		
<b><math>9.2 \pm 1.8 \pm 0.6</math></b>	274	<sup>1,2</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma \pi \pi$	
<sup>1</sup> DOBBS 15 reports $[\Gamma(\psi(2S) \rightarrow \gamma f_0(1500))/\Gamma_{\text{total}}] \times [B(f_0(1500) \rightarrow \pi \pi)] = (3.2 \pm 0.6 \pm 0.2) \times 10^{-5}$ which we divide by our best value $B(f_0(1500) \rightarrow \pi \pi) = (34.9 \pm 2.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					

$\Gamma(\gamma f'_2(1525))/\Gamma_{\text{total}}$					$\Gamma_{142}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	COMMENT		
<b><math>3.3 \pm 0.8 \pm 0.1</math></b>	136	<sup>1,2</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K \bar{K}$	
<sup>1</sup> DOBBS 15 reports $[\Gamma(\psi(2S) \rightarrow \gamma f'_2(1525))/\Gamma_{\text{total}}] \times [B(f'_2(1525) \rightarrow K \bar{K})] = (2.9 \pm 0.6 \pm 0.3) \times 10^{-5}$ which we divide by our best value $B(f'_2(1525) \rightarrow K \bar{K}) = (88.7 \pm 2.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					
<sup>2</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					

$\Gamma(\gamma f_0(1710) \rightarrow \gamma \pi \pi)/\Gamma_{\text{total}}$					$\Gamma_{144}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.5 \pm 0.6</math> OUR AVERAGE</b>					
$3.6 \pm 0.4 \pm 0.5$	290	<sup>1</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma \pi \pi$	
$3.01 \pm 0.41 \pm 1.24$	$35.6 \pm 4.8$	<sup>2</sup> BAI	03c BES	$\psi(2S) \rightarrow \gamma \pi^+ \pi^-$	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					
<sup>2</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = 0.305 \pm 0.016$ .					

$\Gamma(\gamma f_0(1710) \rightarrow \gamma K \bar{K})/\Gamma_{\text{total}}$					$\Gamma_{145}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.6 \pm 0.7</math> OUR AVERAGE</b>					
$6.7 \pm 0.6 \pm 0.6$		375	<sup>1</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K \bar{K}$
$6.04 \pm 0.90 \pm 1.32$	$39.6 \pm 5.9$	<sup>2,3</sup> BAI	03c BES		$\psi(2S) \rightarrow \gamma K^+ K^-$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 15.6$	90	$6.8 \pm 3.1$	<sup>2,3</sup> BAI	03c BES	$\psi(2S) \rightarrow \gamma K_S^0 K_S^0$
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					
<sup>2</sup> Includes unknown branching fractions to $K^+ K^-$ or $K_S^0 K_S^0$ . We have multiplied the $K^+ K^-$ result by a factor of 2 and the $K_S^0 K_S^0$ result by a factor of 4 to obtain the $K \bar{K}$ result.					
<sup>3</sup> Normalized to $B(\psi(2S) \rightarrow J/\psi \pi^+ \pi^-) = 0.305 \pm 0.016$ .					

$\Gamma(\gamma f_0(2100) \rightarrow \gamma \pi \pi)/\Gamma_{\text{total}}$					$\Gamma_{146}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	COMMENT		
<b><math>4.8 \pm 0.5 \pm 0.9</math></b>	373	<sup>1</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma \pi \pi$	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					

$\Gamma(\gamma f_0(2200) \rightarrow \gamma K \bar{K})/\Gamma_{\text{total}}$					$\Gamma_{147}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	COMMENT		
<b><math>3.2 \pm 0.6 \pm 0.8</math></b>	207	<sup>1</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K \bar{K}$	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					

$\Gamma(\gamma f_2(2220) \rightarrow \gamma \pi \pi)/\Gamma_{\text{total}}$					$\Gamma_{148}/\Gamma$
VALUE	CL%	DOCUMENT ID	COMMENT		
<b><math>&lt; 5.8 \times 10^{-6}</math></b>	90	<sup>1,2</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma \pi \pi$	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					
<sup>2</sup> For $\Gamma = 20/50$ MeV, the 90% CL upper limits for $\pi^+ \pi^-$ and $\pi^0 \pi^0$ are $3.2/4.3 \times 10^{-6}$ and $2.6/4.0 \times 10^{-6}$ , respectively.					

$\Gamma(\gamma f_2(2220) \rightarrow \gamma K \bar{K})/\Gamma_{\text{total}}$					$\Gamma_{149}/\Gamma$
VALUE	CL%	DOCUMENT ID	COMMENT		
<b><math>&lt; 9.5 \times 10^{-6}</math></b>	90	<sup>1,2</sup> DOBBS	15	$\psi(2S) \rightarrow \gamma K \bar{K}$	
<sup>1</sup> Using CLEO-c data but not authored by the CLEO Collaboration.					
<sup>2</sup> For $\Gamma = 20/50$ MeV, the 90% CL upper limits for $K^+ K^-$ and $K_S^0 K_S^0$ are $2.1/4.3 \times 10^{-6}$ and $3.7/5.5 \times 10^{-6}$ , respectively.					

$\Gamma(\gamma \eta)/\Gamma_{\text{total}}$					$\Gamma_{151}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.38 \pm 0.48 \pm 0.09</math></b>		13	<sup>1</sup> ABLIKIM	10F BES3	$\psi(2S) \rightarrow \gamma \pi^+ \pi^- \pi^0$ ,
$\gamma 3\pi^0$					
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 2$	90	PEDLAR	09	CLE3	$\psi(2S) \rightarrow \gamma X$
$< 90$	90	BAI	98F	BES	$\psi(2S) \rightarrow \pi^+ \pi^- 3\gamma$
$< 200$	90	YAMADA	77	DASP	$e^+ e^- \rightarrow 3\gamma$
<sup>1</sup> Combining the results from $\eta \rightarrow \pi^+ \pi^- \pi^0$ and $\eta \rightarrow 3\pi^0$ decay modes.					

$\Gamma(\gamma \eta \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{152}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>8.71 \pm 1.25 \pm 1.64</math></b>	418	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \eta \pi^+ \pi^-$	

$\Gamma(\gamma \eta(1405) \rightarrow \gamma K \bar{K} \pi)/\Gamma_{\text{total}}$					$\Gamma_{154}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 0.9</math></b>	90	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma K_S^0 K^+ \pi^- + \text{c.c.}$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.3$	90	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma K^+ K^- \pi^0$	
$< 1.2$	90	<sup>1</sup> SCHARRE	80 MRK1	$e^+ e^-$	
<sup>1</sup> Includes unknown branching fraction $\eta(1405) \rightarrow K \bar{K} \pi$ .					

$\Gamma(\gamma \eta(1405) \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{155}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.36 \pm 0.25 \pm 0.05</math></b>	10	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \eta \pi^+ \pi^-$	

$\Gamma(\gamma \eta(1475) \rightarrow K \bar{K} \pi)/\Gamma_{\text{total}}$					$\Gamma_{157}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 1.4</math></b>	90	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma K^+ K^- \pi^0$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.5$	90	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma K_S^0 K^+ \pi^- + \text{c.c.}$	

$\Gamma(\gamma \eta(1475) \rightarrow \eta \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{158}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 0.88</math></b>	90	ABLIKIM	06R BES2	$\psi(2S) \rightarrow \gamma \eta \pi^+ \pi^-$	

$\Gamma(\gamma(2\pi^+ \pi^-))/\Gamma_{\text{total}}$					$\Gamma_{159}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>39.6 \pm 2.8 \pm 5.0</math></b>	583	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$	

$\Gamma(\gamma K^{*0} K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{160}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>37.0 \pm 6.1 \pm 7.2</math></b>	237	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$	

$\Gamma(\gamma K^{*0} \bar{K}^{*0})/\Gamma_{\text{total}}$					$\Gamma_{161}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>24.0 \pm 4.5 \pm 5.0</math></b>	41	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$	

$\Gamma(\gamma K_S^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{162}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>25.6 \pm 3.6 \pm 3.6</math></b>	115	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$	

$\Gamma(\gamma K^+ K^- \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{163}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>19.1 \pm 2.7 \pm 4.3</math></b>	132	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$	

$\Gamma(\gamma p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{164}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.9 \pm 0.5</math> OUR AVERAGE</b>				Error includes scale factor of 2.0.	
$4.18 \pm 0.26 \pm 0.18$	348	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \gamma p \bar{p}$	
$2.9 \pm 0.4 \pm 0.4$	142	ABLIKIM	07D BES2	$e^+ e^- \rightarrow \psi(2S)$	
<sup>1</sup> From a fit of the $p \bar{p}$ mass distribution to a combination of $\gamma f_2(1950)$ , $\gamma f_2(2150)$ , and $\gamma p \bar{p}$ phase space, for $M(p \bar{p}) < 2.85$ GeV, and accounting for backgrounds from $\psi(2S) \rightarrow \pi^0 p \bar{p}$ and continuum.					

$\Gamma(\gamma f_2(1950) \rightarrow \gamma p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{165}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.2 \pm 0.2 \pm 0.1</math></b>	111	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \gamma p \bar{p}$	

<sup>1</sup> From a fit of the  $p \bar{p}$  mass distribution to a combination of  $\gamma f_2(1950)$ ,  $\gamma f_2(2150)$ , and  $\gamma p \bar{p}$  phase space, for  $M(p \bar{p}) < 2.85$  GeV, and accounting for backgrounds from  $\psi(2S) \rightarrow \pi^0 p \bar{p}$  and continuum.

$\Gamma(\gamma f_2(2150) \rightarrow \gamma p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{166}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.72 \pm 0.18 \pm 0.03</math></b>	73	<sup>1</sup> ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \gamma p \bar{p}$	

<sup>1</sup> From a fit of the  $p \bar{p}$  mass distribution to a combination of  $\gamma f_2(1950)$ ,  $\gamma f_2(2150)$ , and  $\gamma p \bar{p}$  phase space, for  $M(p \bar{p}) < 2.85$  GeV, and accounting for backgrounds from  $\psi(2S) \rightarrow \pi^0 p \bar{p}$  and continuum.

$\Gamma(\gamma X(1835) \rightarrow \gamma p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{167}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>4.57 \pm 0.36 \pm 1.77</math></b>		ABLIKIM	12D BES3	$J/\psi \rightarrow \gamma p \bar{p}$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.6$	90	ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \gamma p \bar{p}$	
$< 5.4$	90	ABLIKIM	07D BES	$\psi(2S) \rightarrow \gamma p \bar{p}$	

$\Gamma(\gamma X \rightarrow \gamma p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{168}/\Gamma$
For a narrow resonance in the range $2.2 < M(X) < 2.8$ GeV.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 2</math></b>	90	ALEXANDER	10 CLEO	$\psi(2S) \rightarrow \gamma p \bar{p}$	



See key on page 601

Meson Particle Listings

$\psi(2S)$

$\Gamma(\gamma\pi^+\pi^-\rho\overline{\rho})/\Gamma_{\text{total}}$		$\Gamma_{169}/\Gamma$		
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$2.8 \pm 1.2 \pm 0.7$	17	ABLIKIM	07D BES2	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\gamma 2(\pi^+\pi^-)K^+K^-)/\Gamma_{\text{total}}$		$\Gamma_{170}/\Gamma$		
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$<22$	90	ABLIKIM	07D BES2	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\gamma 3(\pi^+\pi^-))/\Gamma_{\text{total}}$		$\Gamma_{171}/\Gamma$		
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$<17$	90	ABLIKIM	07D BES2	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\gamma K^+K^-K^+K^-)/\Gamma_{\text{total}}$		$\Gamma_{172}/\Gamma$		
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$<4$	90	ABLIKIM	07D BES2	$e^+e^- \rightarrow \psi(2S)$

$\Gamma(\gamma\gamma J/\psi)/\Gamma_{\text{total}}$		$\Gamma_{173}/\Gamma$		
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$3.1 \pm 0.6 \pm 0.8$ $-1.0$	1.1k	ABLIKIM	12o BES3	$e^+e^- \rightarrow \psi(2S)$

OTHER DECAYS

$\Gamma(\text{invisible})/\Gamma(e^+e^-)$		$\Gamma_{174}/\Gamma_6$		
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.0$	90	LEES	13i BABR	$B \rightarrow K^{(*)}\psi(2S)$

$\psi(2S)$  CROSS-PARTICLE BRANCHING RATIOS

For measurements involving  $B(\psi(2S) \rightarrow \gamma\chi_{cJ}(1P)) \times B(\chi_{cJ}(1P) \rightarrow X)$  see the corresponding entries in the  $\chi_{cJ}(1P)$  sections.

MULTIPOLE AMPLITUDE RATIOS IN RADIATIVE DECAYS  
 $\psi(2S) \rightarrow \gamma\chi_{cJ}(1P)$  and  $\chi_{cJ} \rightarrow \gamma J/\psi(1S)$

$a_2(\chi_{c1})/a_2(\chi_{c2})$  Magnetic quadrupole transition amplitude ratio

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$67^{+19}_{-13}$	59k	<sup>1</sup> ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma\gamma e^+e^-$

<sup>1</sup> Statistical and systematic errors combined. Using values from fits with floating  $M2$  amplitudes  $a_2(\chi_{c1})$ ,  $a_2(\chi_{c2})$ ,  $b_2(\chi_{c1})$ ,  $b_2(\chi_{c2})$  and fixed  $E3$  amplitudes of  $a_3(\chi_{c2}) = b_3(\chi_{c2}) = 0$ . Not independent of values for  $a_2(\chi_{c1}(1P))$  and  $a_2(\chi_{c2}(1P))$  from ARTUSO 09.

$b_2(\chi_{c2})/b_2(\chi_{c1})$  Magnetic quadrupole transition amplitude ratio

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$37^{+53}_{-47}$	59k	<sup>1</sup> ARTUSO	09 CLEO	$\psi(2S) \rightarrow \gamma\gamma e^+e^-$

<sup>1</sup> Statistical and systematic errors combined. Using values from fits with floating  $M2$  amplitudes  $a_2(\chi_{c1})$ ,  $a_2(\chi_{c2})$ ,  $b_2(\chi_{c1})$ ,  $b_2(\chi_{c2})$  and fixed  $E3$  amplitudes of  $a_3(\chi_{c2}) = b_3(\chi_{c2}) = 0$ . Not independent of values for  $a_2(\chi_{c1}(1P))$  and  $b_2(\chi_{c2}(1P))$  from ARTUSO 09.

$\psi(2S)$  REFERENCES

ABLIKIM	15i	PR D91 092006	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	15V	PL B749 414	M. Ablikim et al.	(BES III Collab.)
ANASHIN	15	PL B749 50	V.V. Anashin et al.	(KEDR Collab.)
DOBBS	15	PR D91 052006	S. Dobbs et al.	(NWES)
LEES	15J	PR D92 072008	J.P. Lees et al.	(BABAR Collab.)
ABLIKIM	14G	PR D89 112006	M. Ablikim et al.	(BES III Collab.)
DOBBS	14	PL B739 90	S. Dobbs et al.	(NWES, WAYN)
ABLIKIM	13A	PRL 110 022001	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13D	PR D87 012007	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13F	PR D87 052007	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13M	PR D87 092006	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13R	PR D88 032007	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13S	PR D88 032010	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	13W	PR D88 112007	M. Ablikim et al.	(BES III Collab.)
LEES	13i	PR D87 112005	J.P. Lees et al.	(BABAR Collab.)
LEES	13O	PR D87 092005	J.P. Lees et al.	(BABAR Collab.)
LEES	13Q	PR D88 032013	J.P. Lees et al.	(BABAR Collab.)
LEES	13Y	PR D88 072009	J.P. Lees et al.	(BABAR Collab.)
AAUJ	12H	EPJ C72 1972	R. Aaij et al.	(LHCb Collab.)
ABLIKIM	12D	PRL 108 112003	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12G	PRL 109 042003	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12H	PL B710 594	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12L	PR D86 072011	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12M	PR D86 092008	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12O	PRL 109 172002	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	12Q	CPC 36 1040	M. Ablikim et al.	(BES II Collab.)
ANASHIN	12	PL B711 280	V.V. Anashin et al.	(KEDR Collab.)
LEES	12E	PR D85 112009	J.P. Lees et al.	(BABAR Collab.)
METREVELI	12	PR D85 092007	Z. Metreveli et al.	(NWES, FLOR, WAYN+)
GE	11	PR D84 032008	J.Y. Ge et al.	(CLEO Collab.)
ABLIKIM	10B	PRL 104 132002	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	10F	PRL 105 261801	M. Ablikim et al.	(BES III Collab.)
ALEXANDER	10	PR D82 092002	J.P. Alexander et al.	(CLEO Collab.)
CRONIN-HEN.	10	PR D81 052002	D. Cronin-Hennessey et al.	(CLEO Collab.)
ADAMS	09	PR D80 051106	G.S. Adams et al.	(CLEO Collab.)
ARTUSO	09	PR D80 112003	M. Artuso et al.	(CLEO Collab.)
LIBBY	09	PR D80 072002	J. Libby et al.	(CLEO Collab.)
MITCHELL	09	PRL 102 011801	R.E. Mitchell et al.	(CLEO Collab.)
PEDLAR	09	PR D79 111101	T.K. Pedlar et al.	(CLEO Collab.)
ABLIKIM	08B	PL B659 74	M. Ablikim et al.	(BES Collab.)
ABLIKIM	08C	PL B659 789	M. Ablikim et al.	(BES Collab.)
DOBBS	08A	PRL 101 182003	S. Dobbs et al.	(CLEO Collab.)

MENDEZ	08	PR D78 011102	H. Mendez et al.	(CLEO Collab.)
PDG	08	PL B667 1	C. Amsler et al.	(PDG Collab.)
ABLIKIM	07C	PL B648 149	M. Ablikim et al.	(BES Collab.)
ABLIKIM	07D	PRL 99 011802	M. Ablikim et al.	(BES II Collab.)
ABLIKIM	07H	PR D76 092003	M. Ablikim et al.	(BES Collab.)
ANASHIN	07	JETPL 85 347	V.V. Anashin et al.	(KEDR Collab.)
ANDREOTTI	07	Translated from ZETFP 85 429.	M. Andreotti et al.	(Femilab E835 Collab.)
AUBERT	07AK	PR D76 012008	B. Aubert et al.	(BABAR Collab.)
AUBERT	07AU	PR D76 092005	B. Aubert et al.	(BABAR Collab.)
Also		PR D77 11902E (err.)	B. Aubert et al.	(BABAR Collab.)
AUBERT	07BD	PR D76 092006	B. Aubert et al.	(BABAR Collab.)
PDG	07	Unofficial 2007 WWW edition		(PDG Collab.)
PEDLAR	07	PR D75 011102	T.K. Pedlar et al.	(CLEO Collab.)
ABLIKIM	06G	PR D73 052004	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06I	PR D74 012004	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06L	PRL 97 121801	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06R	PR D74 072001	M. Ablikim et al.	(BES Collab.)
ABLIKIM	06W	PR D74 112003	M. Ablikim et al.	(BES Collab.)
ADAM	06	PRL 96 082004	N.E. Adam et al.	(CLEO Collab.)
AUBERT	06B	PR D73 012005	B. Aubert et al.	(BABAR Collab.)
AUBERT	06D	PR D73 052003	B. Aubert et al.	(BABAR Collab.)
AUBERT,BE	06D	PR D74 091103	B. Aubert et al.	(BABAR Collab.)
DOBBS	06A	PR D74 011105	S. Dobbs et al.	(CLEO Collab.)
ABLIKIM	05E	PR D71 072006	M. Ablikim et al.	(BES Collab.)
ABLIKIM	05H	PR D72 012002	M. Ablikim et al.	(BES Collab.)
ABLIKIM	05I	PL B614 37	M. Ablikim et al.	(BES Collab.)
ABLIKIM	05J	PL B619 247	M. Ablikim et al.	(BES Collab.)
ABLIKIM	05O	PL B630 21	M. Ablikim et al.	(BES Collab.)
ADAM	05	PRL 94 012005	N.E. Adam et al.	(CLEO Collab.)
ADAM	05A	PRL 94 232002	N.E. Adam et al.	(CLEO Collab.)
ANDREOTTI	05	PR D71 032006	M. Andreotti et al.	(FNAL E835 Collab.)
AUBERT	05D	PR D71 052001	B. Aubert et al.	(BABAR Collab.)
BRIERE	05	PRL 95 062001	R.A. Briere et al.	(CLEO Collab.)
PEDLAR	05	PR D72 051108	T.K. Pedlar et al.	(CLEO Collab.)
ROSNER	05	PRL 95 102003	J.L. Rosner et al.	(CLEO Collab.)
ABLIKIM	04B	PR D70 012003	M. Ablikim et al.	(BES Collab.)
ABLIKIM	04K	PR D70 112003	M. Ablikim et al.	(BES Collab.)
ABLIKIM	04L	PR D70 112007	M. Ablikim et al.	(BES Collab.)
ATHAR	04	PR D70 112002	S.B. Athar et al.	(CLEO Collab.)
BAI	04B	PRL 92 052001	J.Z. Bai et al.	(BES Collab.)
BAI	04C	PR D69 072001	J.Z. Bai et al.	(BES Collab.)
BAI	04D	PL B589 7	J.Z. Bai et al.	(BES Collab.)
BAI	04G	PR D70 012004	J.Z. Bai et al.	(BES Collab.)
BAI	04I	PR D70 012006	J.Z. Bai et al.	(BES Collab.)
PDG	04	PL B592 1	S. Eidelman et al.	(PDG Collab.)
SETH	04	PR D69 097503	K.K. Seth	
AULCHENKO	03	PL B573 63	V.M. Aulchenko et al.	(KEDR Collab.)
BAI	03B	PR D67 052002	J.Z. Bai et al.	(BES Collab.)
BAI	03C	PR D67 032004	J.Z. Bai et al.	(BES Collab.)
AUBERT	02B	PR D65 031101	B. Aubert et al.	(BABAR Collab.)
BAI	02	PR D65 052004	J.Z. Bai et al.	(BES Collab.)
BAI	02B	PL B550 24	J.Z. Bai et al.	(BES Collab.)
BAI	02C	PRL 88 101802	J.Z. Bai et al.	(BES Collab.)
PDG	02	PR D66 010001	K. Haglwar et al.	(PDG Collab.)
BAI	01	PR D63 032002	J.Z. Bai et al.	(BES Collab.)
AMBROGIANI	00A	PR D62 032004	M. Ambrogiani et al.	(FNAL E835 Collab.)
ARTAMONOV	00	PL B474 427	A.S. Artamonov et al.	
BAI	00	PRL 84 594	J.Z. Bai et al.	(BES Collab.)
BAI	99C	PRL 83 1918	J.Z. Bai et al.	(BES Collab.)
BAI	98F	PR D57 3854	J.Z. Bai et al.	(BES Collab.)
BAI	98E	PR D58 097101	J.Z. Bai et al.	(BES Collab.)
BAI	98J	PRL 81 5080	J.Z. Bai et al.	(BES Collab.)
ARMSTRONG	97	PR D55 1153	T.A. Armstrong et al.	(E760 Collab.)
GRIBUSHIN	96	PR D53 4723	A. Gribushin et al.	(E706 Collab.)
ARMSTRONG	93B	PR D47 772	T.A. Armstrong et al.	(FNAL E760 Collab.)
ALEXANDER	89	NP B320 45	J.P. Alexander et al.	(LBL, MICH, SLAC)
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RISC, NBS)
GAISER	86	PR D34 711	J. Gaiser et al.	(Crystal Ball Collab.)
KURAEV	85	SJNP 41 466	E.A. Kuraev, V.S. Fadin	(NOVO)
FRANKLIN	83	Translated from YAF 41 733.	M.E.B. Franklin et al.	(LBL, SLAC)
EDWARDS	82C	PRL 48 701	C. Edwards et al.	(CIT, HARV, PRIN+)
LEMOIGNE	82	PL 113B 509	Y. Lemoigne et al.	(SACL, LOIC, SHMP+)
HIMEL	80	PRL 44 920	T. Himel et al.	(LBL, SLAC)
OREGLIA	80	PRL 45 959	M.J. Oreglia et al.	(SLAC, CIT, HARV+)
SCHARRE	80	PL 97B 329	D.L. Scharre et al.	(SLAC, LBL)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents et al.	(NOVO)
Also		SJNP 34 814	A.A. Zholents et al.	(NOVO)
BRANDELIK	79B	NP B160 426	R. Brandelik et al.	(DASP Collab.)
BRANDELIK	79C	ZPHY CL 233	R. Brandelik et al.	(DASP Collab.)
BARTEL	78B	PL 79B 492	W. Bartel et al.	(DESY, HEIDP)
TANENBAUM	78	PR D17 1731	W.M. Tanenbaum et al.	(SLAC, LBL)
BIDDICK	77	PRL 38 1324	C.J. Biddick et al.	(UCSD, UMD, PAVI+)
BRUNSCHE...	77	PL 67B 249	W. Braunschweig et al.	(DASP Collab.)
BURMESTER	77	PL 66B 395	J. Burmester et al.	(DESY, HAMB, SIEG+)
FELDMAN	77	PRPL 33C 285	G.J. Feldman, M.L. Perl	(LBL, SLAC)
YAMADA	77	Hamburg Conf. 69	S. Yamada	(DASP Collab.)
BARTEL	76	PL 64B 483	W. Bartel et al.	(DESY, HEIDP)
TANENBAUM	76	PRL 36 402	W.M. Tanenbaum et al.	(SLAC, LBL) IG
WHITAKER	76	PRL 37 1596	J.S. Whitaker et al.	(SLAC, LBL)
ABRAMS	75	Stanford Symp. 25	G.S. Abrams	(LBL)
ABRAMS	75B	PRL 34 1181	G.S. Abrams et al.	(LBL, SLAC)
BOYARSKI	75C	Palermo Conf. 54	A.M. Boyarski et al.	(SLAC, LBL)
HILGER	75	PRL 35 625	E. Hilger et al.	(STAN, PENN)
LIBERMAN	75	Stanford Symp. 55	A.D. Liberman	(STAN)
LUTH	75	PRL 35 1124	V. Luth et al.	(SLAC, LBL) JPC
WIKI	75	Stanford Symp. 69	B.H. Wiik	(DESY)

## Meson Particle Listings

 $\psi(3770)$  $\psi(3770)$ 

$$J^G(J^{PC}) = 0^-(1^{--})$$

 $\psi(3770)$  MASS (MeV)

OUR FIT includes measurements of  $m_{\psi(2S)}$ ,  $m_{\psi(3770)}$ , and  $m_{\psi(3770)} - m_{\psi(2S)}$ .

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3773.13 ± 0.35 OUR FIT</b>		Error includes scale factor of 1.1.		
<b>3778.1 ± 1.2 OUR AVERAGE</b>				
3779.2 <sup>+1.8</sup> <sub>-1.7</sub> <sup>+0.6</sup> <sub>-0.8</sub>	1	ANASHIN	12A KEDR	$e^+e^- \rightarrow D\bar{D}$
3775.5 ± 2.4 ± 0.5	57	AUBERT	08B BABR	$B \rightarrow D\bar{D}K$
3776 ± 5 ± 4	68	BRODZICKA	08 BELL	$B^+ \rightarrow D^0\bar{D}^0K^+$
3778.8 ± 1.9 ± 0.9		AUBERT	07BE BABR	$e^+e^- \rightarrow D\bar{D}\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3772.0 ± 1.9	2,3	ABLIKIM	08D BES2	$e^+e^- \rightarrow$ hadrons
3778.4 ± 3.0 ± 1.3	34	CHISTOV	04 BELL	Sup. by BRODZICKA 08

<sup>1</sup> Taking into account interference between the resonant and non-resonant  $D\bar{D}$  production.

<sup>2</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = 0^\circ$ .

<sup>3</sup> Interference between the resonant and non-resonant  $D\bar{D}$  production not taken into account.

 $m_{\psi(3770)} - m_{\psi(2S)}$ 

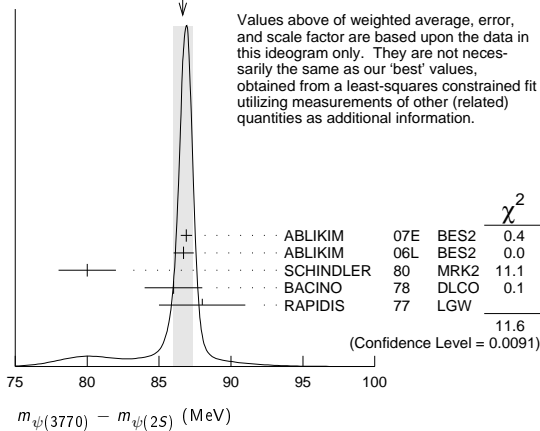
OUR FIT includes measurements of  $m_{\psi(2S)}$ ,  $m_{\psi(3770)}$ , and  $m_{\psi(3770)} - m_{\psi(2S)}$ .

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>87.04 ± 0.35 OUR FIT</b>	Error includes scale factor of 1.1.		
<b>86.6 ± 0.7 OUR AVERAGE</b>	Error includes scale factor of 2.0. See the ideogram below.		
86.9 ± 0.4	4	ABLIKIM	07E BES2 $e^+e^- \rightarrow$ hadrons
86.7 ± 0.7		ABLIKIM	06L BES2 $e^+e^- \rightarrow$ hadrons
80 ± 2		SCHINDLER	80 MRK2 $e^+e^-$
86 ± 2	5	BACINO	78 DLCO $e^+e^-$
88 ± 3		RAPIDIS	77 LGW $e^+e^-$

<sup>4</sup> BES-II  $\psi(2S)$  mass subtracted (see ABLIKIM 06L).

<sup>5</sup> SPEAR  $\psi(2S)$  mass subtracted (see SCHINDLER 80).

WEIGHTED AVERAGE  
86.6 ± 0.7 (Error scaled by 2.0)

 $\psi(3770)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>27.2 ± 1.0 OUR FIT</b>				
<b>27.5 ± 0.9 OUR AVERAGE</b>				
24.9 <sup>+4.6</sup> <sub>-4.0</sub> <sup>+0.5</sup> <sub>-1.1</sub>		6 ANASHIN	12A KEDR	$e^+e^- \rightarrow D\bar{D}$
30.4 ± 8.5	7,8	ABLIKIM	08D BES2	$e^+e^- \rightarrow$ hadrons
27 ± 10 ± 5	68	BRODZICKA	08 BELL	$B^+ \rightarrow D^0\bar{D}^0K^+$
28.5 ± 1.2 ± 0.2		ABLIKIM	07E BES2	$e^+e^- \rightarrow$ hadrons
23.5 ± 3.7 ± 0.9		AUBERT	07BE BABR	$e^+e^- \rightarrow D\bar{D}\gamma$
26.9 ± 2.4 ± 0.3		ABLIKIM	06L BES2	$e^+e^- \rightarrow$ hadrons
24 ± 5		SCHINDLER	80 MRK2	$e^+e^-$
24 ± 5		BACINO	78 DLCO	$e^+e^-$
28 ± 5		RAPIDIS	77 LGW	$e^+e^-$

<sup>6</sup> Taking into account interference between the resonant and non-resonant  $D\bar{D}$  production.

<sup>7</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = 0^\circ$ .

<sup>8</sup> Interference between the resonant and non-resonant  $D\bar{D}$  production not taken into account.

 $\psi(3770)$  DECAY MODES

In addition to the dominant decay mode to  $D\bar{D}$ ,  $\psi(3770)$  was found to decay into the final states containing the  $J/\psi$  (BAI 05, ADAM 06). ADAMS 06 and HUANG 06A searched for various decay modes with light hadrons and found a statistically significant signal for the decay to  $\phi\eta$  only (ADAMS 06).

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor / Confidence level
$\Gamma_1 D\bar{D}$	(93 <sup>+8</sup> <sub>-9</sub> ) %	S=2.0
$\Gamma_2 D^0\bar{D}^0$	(52 <sup>+4</sup> <sub>-5</sub> ) %	S=2.0
$\Gamma_3 D^+D^-$	(41 ± 4) %	S=2.0
$\Gamma_4 J/\psi\pi^+\pi^-$	(1.93 ± 0.28) × 10 <sup>-3</sup>	
$\Gamma_5 J/\psi\pi^0\pi^0$	(8.0 ± 3.0) × 10 <sup>-4</sup>	
$\Gamma_6 J/\psi\eta$	(9 ± 4) × 10 <sup>-4</sup>	
$\Gamma_7 J/\psi\pi^0$	< 2.8 × 10 <sup>-4</sup>	CL=90%
$\Gamma_8 e^+e^-$	(9.6 ± 0.7) × 10 <sup>-6</sup>	S=1.3

## Decays to light hadrons

$\Gamma_9 b_1(1235)\pi$	< 1.4 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{10} \phi\eta'$	< 7 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{11} \omega\eta'$	< 4 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{12} \rho^0\eta'$	< 6 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{13} \phi\eta$	(3.1 ± 0.7) × 10 <sup>-4</sup>	
$\Gamma_{14} \omega\eta$	< 1.4 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{15} \rho^0\eta$	< 5 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{16} \phi\pi^0$	< 3 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{17} \omega\pi^0$	< 6 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{18} \pi^+\pi^-\pi^0$	< 5 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{19} \rho\pi$	< 5 × 10 <sup>-6</sup>	CL=90%
$\Gamma_{20} K^+K^-$		
$\Gamma_{21} K^*(892)^+K^- + c.c.$	< 1.4 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{22} K^*(892)^0\bar{K}^0 + c.c.$	< 1.2 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{23} K_S^0 K_L^0$	< 1.2 × 10 <sup>-5</sup>	CL=90%
$\Gamma_{24} 2(\pi^+\pi^-)$	< 1.12 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{25} 2(\pi^+\pi^-)\pi^0$	< 1.06 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{26} 2(\pi^+\pi^-\pi^0)$	< 5.85 %	CL=90%
$\Gamma_{27} \omega\pi^+\pi^-$	< 6.0 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{28} 3(\pi^+\pi^-)$	< 9.1 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{29} 3(\pi^+\pi^-)\pi^0$	< 1.37 %	CL=90%
$\Gamma_{30} 3(\pi^+\pi^-)2\pi^0$	< 11.74 %	CL=90%
$\Gamma_{31} \eta\pi^+\pi^-$	< 1.24 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{32} \pi^+\pi^-2\pi^0$	< 8.9 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{33} \rho^0\pi^+\pi^-$	< 6.9 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{34} \eta 3\pi$	< 1.34 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{35} \eta 2(\pi^+\pi^-)$	< 2.43 %	CL=90%
$\Gamma_{36} \eta\rho^0\pi^+\pi^-$	< 1.45 %	CL=90%
$\Gamma_{37} \eta' 3\pi$	< 2.44 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{38} K^+K^-\pi^+\pi^-$	< 9.0 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{39} \phi\pi^+\pi^-$	< 4.1 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{40} K^+K^-2\pi^0$	< 4.2 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{41} 4(\pi^+\pi^-)$	< 1.67 %	CL=90%
$\Gamma_{42} 4(\pi^+\pi^-)\pi^0$	< 3.06 %	CL=90%
$\Gamma_{43} \phi f_0(980)$	< 4.5 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{44} K^+K^-\pi^+\pi^-\pi^0$	< 2.36 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{45} K^+K^-\rho^0\pi^0$	< 8 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{46} K^+K^-\rho^+\pi^-$	< 1.46 %	CL=90%
$\Gamma_{47} \omega K^+K^-$	< 3.4 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{48} \phi\pi^+\pi^-\pi^0$	< 3.8 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{49} K^{*0}K^-\pi^+\pi^0 + c.c.$	< 1.62 %	CL=90%
$\Gamma_{50} K^{*+}K^-\pi^+\pi^- + c.c.$	< 3.23 %	CL=90%
$\Gamma_{51} K^+K^-\pi^+\pi^-2\pi^0$	< 2.67 %	CL=90%
$\Gamma_{52} K^+K^-2(\pi^+\pi^-)$	< 1.03 %	CL=90%
$\Gamma_{53} K^+K^-2(\pi^+\pi^-)\pi^0$	< 3.60 %	CL=90%
$\Gamma_{54} \eta K^+K^-$	< 4.1 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{55} \eta K^+K^-\pi^+\pi^-$	< 1.24 %	CL=90%
$\Gamma_{56} \rho^0 K^+K^-$	< 5.0 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{57} 2(K^+K^-)$	< 6.0 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{58} \phi K^+K^-$	< 7.5 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{59} 2(K^+K^-)\pi^0$	< 2.9 × 10 <sup>-4</sup>	CL=90%
$\Gamma_{60} 2(K^+K^-)\pi^+\pi^-$	< 3.2 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{61} K_S^0 K^-\pi^+$	< 3.2 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{62} K_S^0 K^-\pi^+\pi^0$	< 1.33 %	CL=90%
$\Gamma_{63} K_S^0 K^-\rho^+$	< 6.6 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{64} K_S^0 K^-\pi^+\pi^-$	< 8.7 × 10 <sup>-3</sup>	CL=90%

See key on page 601

## Meson Particle Listings

 $\psi(3770)$ 

$\Gamma_{65}$	$K_S^0 K^- \pi^+ \rho^0$	$< 1.6$	%	CL=90%
$\Gamma_{66}$	$K_S^0 K^- \pi^+ \eta$	$< 1.3$	%	CL=90%
$\Gamma_{67}$	$K_S^0 K^- 2\pi^+ \pi^- \pi^0$	$< 4.18$	%	CL=90%
$\Gamma_{68}$	$K_S^0 K^- 2\pi^+ \pi^- \eta$	$< 4.8$	%	CL=90%
$\Gamma_{69}$	$K_S^0 K^- \pi^+ 2(\pi^+ \pi^-)$	$< 1.22$	%	CL=90%
$\Gamma_{70}$	$K_S^0 K^- \pi^+ 2\pi^0$	$< 2.65$	%	CL=90%
$\Gamma_{71}$	$K_S^0 K^- K^+ K^- \pi^+$	$< 4.9$	$\times 10^{-3}$	CL=90%
$\Gamma_{72}$	$K_S^0 K^- K^+ K^- \pi^+ \pi^0$	$< 3.0$	%	CL=90%
$\Gamma_{73}$	$K_S^0 K^- K^+ K^- \pi^+ \eta$	$< 2.2$	%	CL=90%
$\Gamma_{74}$	$K_S^0 K^- \pi^+ + \text{c.c.}$	$< 9.7$	$\times 10^{-3}$	CL=90%
$\Gamma_{75}$	$\rho \bar{\rho}$	$< 4$	$\times 10^{-5}$	CL=90%
$\Gamma_{76}$	$\rho \bar{\rho} \pi^0$	$< 5.8$	$\times 10^{-4}$	CL=90%
$\Gamma_{77}$	$\rho \bar{\rho} \pi^+ \pi^-$	$< 1.2$	$\times 10^{-4}$	CL=90%
$\Gamma_{78}$	$\Lambda \bar{\Lambda}$	$< 1.85$	$\times 10^{-3}$	CL=90%
$\Gamma_{79}$	$\rho \bar{\rho} \pi^+ \pi^- \pi^0$	$< 2.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{80}$	$\omega \rho \bar{\rho}$	$< 7$	$\times 10^{-5}$	CL=90%
$\Gamma_{81}$	$\Lambda \bar{\Lambda} \pi^0$	$< 2.6$	$\times 10^{-3}$	CL=90%
$\Gamma_{82}$	$\rho \bar{\rho} 2(\pi^+ \pi^-)$	$< 5.4$	$\times 10^{-4}$	CL=90%
$\Gamma_{83}$	$\eta \rho \bar{\rho}$	$< 3.3$	$\times 10^{-3}$	CL=90%
$\Gamma_{84}$	$\eta \rho \bar{\rho} \pi^+ \pi^-$	$< 1.7$	$\times 10^{-3}$	CL=90%
$\Gamma_{85}$	$\rho^0 \rho \bar{\rho}$	$< 3.2$	$\times 10^{-4}$	CL=90%
$\Gamma_{86}$	$\rho \bar{\rho} K^+ K^-$	$< 6.9$	$\times 10^{-3}$	CL=90%
$\Gamma_{87}$	$\eta \rho \bar{\rho} K^+ K^-$	$< 1.2$	$\times 10^{-3}$	CL=90%
$\Gamma_{88}$	$\pi^0 \rho \bar{\rho} K^+ K^-$	$< 1.3$	$\times 10^{-4}$	CL=90%
$\Gamma_{89}$	$\phi \rho \bar{\rho}$	$< 2.5$	$\times 10^{-4}$	CL=90%
$\Gamma_{90}$	$\Lambda \bar{\Lambda} \pi^+ \pi^-$	$< 2.8$	$\times 10^{-4}$	CL=90%
$\Gamma_{91}$	$\Lambda \bar{\rho} K^+$	$< 6.3$	$\times 10^{-4}$	CL=90%
$\Gamma_{92}$	$\Lambda \bar{\rho} K^+ \pi^+ \pi^-$	$< 1.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{93}$	$\Lambda \bar{\Lambda} \eta$	$< 1.0$	$\times 10^{-4}$	CL=90%
$\Gamma_{94}$	$\Sigma^+ \bar{\Sigma}^-$	$< 4$	$\times 10^{-5}$	CL=90%
$\Gamma_{95}$	$\Sigma^0 \bar{\Sigma}^0$	$< 1.5$	$\times 10^{-4}$	CL=90%
$\Gamma_{96}$	$\Xi^+ \bar{\Xi}^-$	$< 1.4$	$\times 10^{-4}$	CL=90%
$\Gamma_{97}$	$\Xi^0 \bar{\Xi}^0$			

## Radiative decays

$\Gamma_{98}$	$\gamma \chi_{c2}$	$< 6.4$	$\times 10^{-4}$	CL=90%
$\Gamma_{99}$	$\gamma \chi_{c1}$	$(2.48 \pm 0.23) \times 10^{-3}$		
$\Gamma_{100}$	$\gamma \chi_{c0}$	$(7.0 \pm 0.6) \times 10^{-3}$		
$\Gamma_{101}$	$\gamma \eta_c$	$< 7$	$\times 10^{-4}$	CL=90%
$\Gamma_{102}$	$\gamma \eta_c(2S)$	$< 9$	$\times 10^{-4}$	CL=90%
$\Gamma_{103}$	$\gamma \eta'$	$< 1.8$	$\times 10^{-4}$	CL=90%
$\Gamma_{104}$	$\gamma \eta$	$< 1.5$	$\times 10^{-4}$	CL=90%
$\Gamma_{105}$	$\gamma \pi^0$	$< 2$	$\times 10^{-4}$	CL=90%

## CONSTRAINED FIT INFORMATION

An overall fit to the total width, a partial width, and 3 branching ratios uses 23 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 20.1$  for 19 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	99		
$x_8$	0	0	
$\Gamma$	0	0	-44
	$x_2$	$x_3$	$x_8$

	Mode	Rate (MeV)	Scale factor
$\Gamma_2$	$D^0 \bar{D}^0$	$14.0 \pm 1.4$	1.8
$\Gamma_3$	$D^+ D^-$	$11.2 \pm 1.1$	1.7
$\Gamma_8$	$e^+ e^-$	$(2.62 \pm 0.18) \times 10^{-4}$	1.4

 $\psi(3770)$  PARTIAL WIDTHS

$\Gamma(e^+e^-)$	$\Gamma_8$			
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.262 \pm 0.018</math> OUR FIT</b>	Error	includes scale factor of 1.4.		
<b><math>0.256 \pm 0.016</math> OUR AVERAGE</b>	Error	includes scale factor of 1.2.		
$0.154^{+0.079+0.021}_{-0.058-0.027}$	9,10	ANASHIN	12A	KEDR $e^+e^- \rightarrow D\overline{D}$
$0.22 \pm 0.05$	11,12	ABLIKIM	08D	BES2 $e^+e^- \rightarrow$ hadrons
$0.277 \pm 0.011 \pm 0.013$	12	ABLIKIM	07E	BES2 $e^+e^- \rightarrow$ hadrons

$0.203 \pm 0.003 \pm 0.041$	1.4M	12,13	BESSION	06	CLEO	$e^+ e^- \rightarrow \text{hadrons}$
$0.276 \pm 0.050$		12	SCHINDLER	80	MRK2	$e^+ e^-$
$0.18 \pm 0.06$		12	BACINO	78	DLCO	$e^+ e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.414 \pm 0.072 \pm 0.093$		10,14	ANASHIN	12A	KEDR	$e^+ e^- \rightarrow D \bar{D}$
$-0.080 - 0.028$		15	RAPIDIS	77	LGW	$e^+ e^-$
$0.37 \pm 0.09$						
9 Solution I of the two solutions.						
10 Taking into account interference between the resonant and non-resonant $D \bar{D}$ production.						
11 Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the $\psi(3770)$ , $\psi(4040)$ , $\psi(4160)$ , and $\psi(4415)$ resonances. Phase angle fixed in the fit to $\delta = 0^\circ$ .						
12 Interference between the resonant and non-resonant $D \bar{D}$ production not taken into account.						
13 BESSION 06 (as corrected in BESSION 10) measure $\sigma(e^+ e^- \rightarrow \psi(3770) \rightarrow \text{hadrons}) = 6.36 \pm 0.08 \pm 0.41$ nb at $\sqrt{s} = 3773 \pm 1$ MeV, and obtain $\Gamma_{ee}$ from the Born-level cross section calculated using $\psi(3770)$ mass and width from our 2004 edition, PDG 04.						
14 Solution II of the two solutions.						
15 See also $\Gamma(e^+ e^-) / \Gamma_{\text{total}}$ below.						

 $\psi(3770)$  BRANCHING RATIOS

$\Gamma(D\bar{D})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma = (\Gamma_2+\Gamma_3)/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.93^{+0.08}_{-0.09}</math> OUR FIT</b>	Error includes scale factor of 2.0.			
<b><math>0.93^{+0.08}_{-0.09}</math> OUR AVERAGE</b>	Error includes scale factor of 2.1.			
$0.849 \pm 0.056 \pm 0.018$	$^{16}$	ABLIKIM	08B	BES2 $e^+e^- \rightarrow \text{non-}D\bar{D}$
$1.033 \pm 0.014^{+0.048}_{-0.066}$	$1.427\text{M}$ $^{17}$	BESSION	06	CLEO $e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.866 \pm 0.050 \pm 0.036$	$^{18,19}$	ABLIKIM	07K	BES2 $e^+e^- \rightarrow \text{non-}D\bar{D}$
$0.836 \pm 0.073 \pm 0.042$	$^{19}$	ABLIKIM	06L	BES2 $e^+e^- \rightarrow D\bar{D}$
$0.855 \pm 0.017 \pm 0.058$	$^{19,20}$	ABLIKIM	06N	BES2 $e^+e^- \rightarrow D\bar{D}$

$\Gamma(D^0\bar{D}^0)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.52 <math>\pm_{-0.05}^{+0.04}</math> OUR FIT</b>	Error includes scale factor of 2.0.			
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.467 $\pm 0.047 \pm 0.023$	ABLIKIM	06L	BES2	$e^+e^- \rightarrow D^0\bar{D}^0$
0.499 $\pm 0.013 \pm 0.038$	20 ABLIKIM	06N	BES2	$e^+e^- \rightarrow D^0\bar{D}^0$

$\Gamma(D^+ D^-)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.41 ±0.04 OUR FIT</b>	Error includes scale factor of 2.0.			
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.369±0.037±0.028	20	ABLIKIM	06L	BES2 $e^+e^- \rightarrow D^+D^-$
0.357±0.011±0.034	20	ABLIKIM	06N	BES2 $e^+e^- \rightarrow D^+D^-$

$\Gamma(D^0\bar{D}^0)/\Gamma(D^+D^-)$				$\Gamma_2/\Gamma_3$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.253±0.016 OUR FIT</b>				
<b>1.253±0.016 OUR AVERAGE</b>				
1.252±0.009±0.013	5.3M	BONVICINI	14	CLEO $e^+e^- \rightarrow D\bar{D}$
1.39±0.31±0.12		PAKHLOVA	08	BELL $10.6 e^+e^- \rightarrow D\bar{D}\gamma$
1.78±0.33±0.24		AUBERT	07BE	BABR $e^+e^- \rightarrow D\bar{D}\gamma$
1.27±0.12±0.08		ABLIKIM	06L	BES2 $e^+e^- \rightarrow D\bar{D}$
2.43±1.50±0.43	34	21 CHISTOV	04	BELL $B^+ \rightarrow \psi(3770) K^+$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1.258±0.016±0.014	22	DOBBS	07	CLEO $e^+e^- \rightarrow D\bar{D}$

$\Gamma(J/\psi\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.93±0.28 OUR AVERAGE</b>				
1.89±0.20±0.20	231 ± 33	ADAM	06	CLEO $e^+e^- \rightarrow \psi(3770)$
3.4 ±1.4 ±0.9	17.8 ± 4.8	BAI	05	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(J/\psi\pi^0\pi^0)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.080 \pm 0.025 \pm 0.016</math></b>	$39 \pm 14$	ADAM	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(J/\psi\eta)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>87 \pm 33 \pm 22</math></b>	$22 \pm 10$	ADAM	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(J/\psi\pi^0)/\Gamma_{\text{total}}$			$\Gamma_7/\Gamma$		
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<28	90	<10	ADAM	06	CLEO $e^+e^- \rightarrow \psi(3770)$

# Meson Particle Listings

## $\psi(3770)$

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$			
VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT	
<b><math>0.96 \pm 0.07</math> OUR FIT</b>	Error includes scale factor of 1.3.			
<b><math>1.3 \pm 0.2</math></b>	RAPIDIS	77	LGW	$e^+e^-$
<sup>16</sup> Neglecting interference.				
<sup>17</sup> Obtained by comparing a measurement of the total cross section (corrected in BESSON 10) with that of $D\bar{D}$ reported by CLEO in DOBBS 07.				
<sup>18</sup> Using $\sigma^{obs} = 7.07 \pm 0.58$ nb and neglecting interference.				
<sup>19</sup> Not independent of ABLIKIM 08b.				
<sup>20</sup> From a measurement of $\sigma(e^+e^- \rightarrow D\bar{D})$ at $\sqrt{s} = 3773$ MeV, using the $\psi(3770)$ resonance parameters measured by ABLIKIM 06L.				
<sup>21</sup> See ADLER 88c for older measurements of this quantity.				
<sup>22</sup> Superseded by BONVICINI 14.				

### DECAYS TO LIGHT HADRONS

$\Gamma(b_1(1235)\pi)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.4</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\phi\eta')/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;7</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\omega\eta')/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;4</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\rho^0\eta')/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;6</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\phi\eta)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>3.1 \pm 0.6 \pm 0.3</math></b>		<sup>23</sup> ADAMS	06	CLEO $3.773\ e^+e^- \rightarrow \phi\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<19$	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\omega\eta)/\Gamma_{\text{total}}$	$\Gamma_{14}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.4</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\rho^0\eta)/\Gamma_{\text{total}}$	$\Gamma_{15}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\phi\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{16}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<50$	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{17}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;6</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{18}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5</math></b>	90	<sup>23,25</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\rho\pi)/\Gamma_{\text{total}}$	$\Gamma_{19}/\Gamma$			
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5</math></b>	90	<sup>23,25</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(K^+K^-)/\Gamma_{\text{total}}$	$\Gamma_{20}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 10^{-5}$	<sup>26</sup> DRUZHININ	15	RVUE	$e^+e^- \rightarrow \psi(3770)$

$\Gamma(K^*(892)^+K^- + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{21}/\Gamma$			
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.4</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(K^*(892)^0\bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{22}/\Gamma$			
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.2</math></b>	90	<sup>23</sup> ADAMS	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(K_S^0K_L^0)/\Gamma_{\text{total}}$	$\Gamma_{23}/\Gamma$			
VALUE (units $10^{-9}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.2</math></b>	90	<sup>27</sup> CRONIN-HEN..06	CLEO	$e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<21$	90	<sup>28</sup> ABLIKIM	04f	BES $e^+e^- \rightarrow \psi(3770)$

$\Gamma(2(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_{24}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;11.2</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<48$	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(2(\pi^+\pi^-\pi^0))/\Gamma_{\text{total}}$	$\Gamma_{25}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;10.6</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<62$	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(2(\pi^+\pi^-\pi^0))/\Gamma_{\text{total}}$	$\Gamma_{26}/\Gamma$			
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN
<b><math>&lt;58.5</math></b>	90	305	ABLIKIM	08N
			BES2	$e^+e^- \rightarrow \psi(3770)$

$\Gamma(\omega\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{27}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;6.0</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<55$	90	<sup>24</sup> ABLIKIM	07i	BES2 $3.77\ e^+e^-$

$\Gamma(3(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_{28}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;91</math></b>	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(3(\pi^+\pi^-\pi^0))/\Gamma_{\text{total}}$	$\Gamma_{29}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;137</math></b>	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(3(\pi^+\pi^-2\pi^0))/\Gamma_{\text{total}}$	$\Gamma_{30}/\Gamma$			
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN
<b><math>&lt;117.4</math></b>	90	59	ABLIKIM	08N
			BES2	$e^+e^- \rightarrow \psi(3770)$

$\Gamma(\eta\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{31}/\Gamma$			
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.24</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.3$	90	<sup>24</sup> ABLIKIM	10D	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{32}/\Gamma$			
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN
<b><math>&lt;8.9</math></b>	90	218	ABLIKIM	08N
			BES2	$e^+e^- \rightarrow \psi(3770)$

$\Gamma(\rho^0\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{33}/\Gamma$			
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;6.9</math></b>	90	<sup>24</sup> ABLIKIM	07f	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\eta3\pi)/\Gamma_{\text{total}}$	$\Gamma_{34}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;13.4</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\eta2(\pi^+\pi^-))/\Gamma_{\text{total}}$	$\Gamma_{35}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;243</math></b>	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\eta\rho^0\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{36}/\Gamma$			
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.45</math></b>	90	<sup>24</sup> ABLIKIM	10D	BES2 $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\eta'3\pi)/\Gamma_{\text{total}}$	$\Gamma_{37}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;24.4</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(K^+K^-\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{38}/\Gamma$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;9.0</math></b>	90	<sup>29</sup> HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<48$	90	<sup>24</sup> ABLIKIM	07b	BES2 $e^+e^- \rightarrow \psi(3770)$

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Meson Particle Listings

$\psi(3770)$

$\Gamma(\phi\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{39}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 4.1	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<16	90	24	ABLIKIM	07B	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K^+K^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{40}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<4.2	90	14	ABLIKIM	08N	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(4(\pi^+\pi^-))/\Gamma_{\text{total}}$					$\Gamma_{41}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<16.7	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(4(\pi^+\pi^-\pi^0))/\Gamma_{\text{total}}$					$\Gamma_{42}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<30.6	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(\phi f_0(980))/\Gamma_{\text{total}}$					$\Gamma_{43}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<4.5	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K^+K^-\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{44}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 23.6	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<111	90	24	ABLIKIM	07B	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K^+K^-\rho^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{45}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<8	90	24	ABLIKIM	07I	BES2 3.77 $e^+e^-$				
$\Gamma(K^+K^-\rho^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{46}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<146	90	24	ABLIKIM	07I	BES2 3.77 $e^+e^-$				
$\Gamma(\omega K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{47}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 3.4	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<66	90	24	ABLIKIM	07I	BES2 3.77 $e^+e^-$				
$\Gamma(\phi\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{48}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<38	90	24	ABLIKIM	07I	BES2 3.77 $e^+e^-$				
$\Gamma(K^{*0}K^-\pi^+\pi^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{49}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<162	90	24	ABLIKIM	07I	BES2 3.77 $e^+e^-$				
$\Gamma(K^{*+}K^-\pi^+\pi^- + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<323	90	24	ABLIKIM	07I	BES2 3.77 $e^+e^-$				
$\Gamma(K^+K^-\pi^+\pi^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{51}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<26.7	90	24	ABLIKIM	08N	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K^+K^-2(\pi^+\pi^-))/\Gamma_{\text{total}}$					$\Gamma_{52}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<10.3	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K^+K^-2(\pi^+\pi^-)\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{53}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<36.0	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(\eta K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{54}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 4.1	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<31	90	24	ABLIKIM	10D	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(\eta K^+K^-\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{55}/\Gamma$				
VALUE (units $10^{-2}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<1.24	90	24	ABLIKIM	10D	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(\rho^0 K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{56}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<5.0	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				

$\Gamma(2(K^+K^-))/\Gamma_{\text{total}}$					$\Gamma_{57}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 6.0	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<17	90	24	ABLIKIM	07B	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(\phi K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{58}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 7.5	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<24	90	24	ABLIKIM	07B	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(2(K^+K^-)\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{59}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
< 2.9	90	29	HUANG	06A	CLEO $e^+e^- \rightarrow \psi(3770)$				
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<46	90	24	ABLIKIM	07B	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(2(K^+K^-)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{60}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<3.2	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{61}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<3.2	90	18	ABLIKIM	08M	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\pi^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{62}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<13.3	90	40	ABLIKIM	08M	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\rho^+)/\Gamma_{\text{total}}$					$\Gamma_{63}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<6.6	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-2\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{64}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<8.7	90	39	ABLIKIM	08M	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\pi^+\rho^0)/\Gamma_{\text{total}}$					$\Gamma_{65}/\Gamma$				
VALUE (units $10^{-2}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<1.6	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\pi^+\eta)/\Gamma_{\text{total}}$					$\Gamma_{66}/\Gamma$				
VALUE (units $10^{-2}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<1.3	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-2\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{67}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<41.8	90	23	ABLIKIM	08M	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-2\pi^+\pi^-\eta)/\Gamma_{\text{total}}$					$\Gamma_{68}/\Gamma$				
VALUE (units $10^{-2}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<4.8	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\pi^+2(\pi^+\pi^-))/\Gamma_{\text{total}}$					$\Gamma_{69}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<12.2	90	4	ABLIKIM	08M	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-\pi^+2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{70}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT				
<26.5	90	17	ABLIKIM	08M	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-K^+K^-\pi^+)/\Gamma_{\text{total}}$					$\Gamma_{71}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<4.9	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-K^+K^-\pi^+\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{72}/\Gamma$				
VALUE (units $10^{-2}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<3.0	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K_S^0 K^-K^+K^-\pi^+\eta)/\Gamma_{\text{total}}$					$\Gamma_{73}/\Gamma$				
VALUE (units $10^{-2}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<2.2	90		ABLIKIM	09C	BES2 $e^+e^- \rightarrow \psi(3770)$				
$\Gamma(K^{*0} K^-\pi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{74}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%		DOCUMENT ID	TECN	COMMENT				
<9.7	90	24	ABLIKIM	07F	BES2 $e^+e^- \rightarrow \psi(3770)$				

# Meson Particle Listings

## $\psi(3770)$

$\Gamma(p\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{75}/\Gamma$				
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •									
$7.1^{+8.6}_{-2.9}$	90	30 ABLIKIM	14L BES3	$e^+e^- \rightarrow \psi(3770)$					
$310 \pm 30$	90	31 ABLIKIM	14L BES3	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(p\overline{p}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{76}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 0.4	90	32,33 ABLIKIM	14o BES3	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
$59^{+3}_{-2} \pm 5$	90	32,34 ABLIKIM	14o BES3	$e^+e^- \rightarrow \psi(3770)$					
<12	90	24 ABLIKIM	07B BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(p\overline{p}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{77}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 5.8	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<16	90	24 ABLIKIM	07B BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Lambda\overline{\Lambda})/\Gamma_{\text{total}}$					$\Gamma_{78}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.2	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<4	90	24 ABLIKIM	07F BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(p\overline{p}\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{79}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<18.5	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<73	90	24 ABLIKIM	07B BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\omega p\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{80}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 2.9	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<30	90	35 ABLIKIM	07I BES2	$3.77 e^+e^-$					

$\Gamma(\Lambda\overline{\Lambda}\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{81}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 0.7	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<12	90	24 ABLIKIM	07I BES2	$3.77 e^+e^-$					

$\Gamma(p\overline{p}2(\pi^+\pi^-))/\Gamma_{\text{total}}$					$\Gamma_{82}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.6	90	24 ABLIKIM	07F BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\eta p\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{83}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 5.4	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<11	90	24 ABLIKIM	10D BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\eta p\overline{p}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{84}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<3.3	90	24 ABLIKIM	10D BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\rho^0 p\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{85}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.7	90	24 ABLIKIM	07F BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(p\overline{p}K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{86}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 3.2	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<11	90	24 ABLIKIM	07B BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\eta p\overline{p}K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{87}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<6.9	90	24 ABLIKIM	10D BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\pi^0 p\overline{p}K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{88}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.2	90	24 ABLIKIM	10D BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\phi p\overline{p})/\Gamma_{\text{total}}$					$\Gamma_{89}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.3	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<9	90	24 ABLIKIM	07B BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Lambda\overline{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{90}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 2.5	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
< 4.7	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					
<39	90	24 ABLIKIM	07F BES2	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Lambda\overline{\Lambda}K^+)/\Gamma_{\text{total}}$					$\Gamma_{91}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.8	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Lambda\overline{\Lambda}K^+\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{92}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<6.3	90	29 HUANG	06A CLEO	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Lambda\overline{\Lambda}\eta)/\Gamma_{\text{total}}$					$\Gamma_{93}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.9	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Sigma^+\Sigma^-)/\Gamma_{\text{total}}$					$\Gamma_{94}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.0	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Sigma^0\Sigma^0)/\Gamma_{\text{total}}$					$\Gamma_{95}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<0.4	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Xi^+\Xi^-)/\Gamma_{\text{total}}$					$\Gamma_{96}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.5	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					

$\Gamma(\Xi^0\Xi^0)/\Gamma_{\text{total}}$					$\Gamma_{97}/\Gamma$				
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.4	90	36 ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(3770)$					

- 23 Comparing cross sections at  $\sqrt{s} = 3.773$  GeV and  $\sqrt{s} = 3.671$  GeV, neglecting interference, and using  $\sigma(\psi(3770) \rightarrow D\overline{D}) = 6.39 \pm 0.20$  nb.
- 24 Assuming that interference effects between resonance and continuum can be neglected and using  $\sigma^{obs}(e^+e^- \rightarrow \psi(3770)) = 7.15 \pm 0.38$  nb.
- 25 Data suggest possible destructive interference with continuum.
- 26 DRUZHININ 15 uses BABAR and CLEO data takitaking into account interference of the processes  $e^+e^- \rightarrow K^+K^-$  and  $e^+e^- \rightarrow K_S^0\overline{K}_L^0$ .
- 27 Using  $\sigma(e^+e^- \rightarrow \psi(3770) \rightarrow \text{hadrons}) = (6.38 \pm 0.08^{+0.41}_{-0.30})$  nb from BESSON 06 and  $B(K_S^0 \rightarrow \pi^+\pi^-) = 0.6895 \pm 0.0014$ .
- 28 Using  $B(K_S^0 \rightarrow \pi^+\pi^-) = 0.6860 \pm 0.0027$ .
- 29 Using  $\sigma_{tot}(e^+e^- \rightarrow \psi(3770)) = 7.9 \pm 0.6$  nb at the resonance.
- 30 Solution I of two equivalent solutions in a fit with a resonance interfering with continuum.
- 31 Solution II of two equivalent solutions in a fit with a resonance interfering with continuum.
- 32 Calculated by the authors using  $\sigma(e^+e^- \rightarrow \psi(3770) \rightarrow \text{hadrons}) = 6.36 \pm 0.08^{+0.41}_{-0.30}$  nb from BESSON 10.
- 33 Solution I of two equivalent solutions in a fit with a resonance interfering with continuum.
- 34 Solution II of two equivalent solutions in a fit with a resonance interfering with continuum.
- 35 Using  $\sigma^{obs} = 7.15 \pm 0.27 \pm 0.27$  nb and neglecting interference.
- 36 Assuming that interference effects between resonance and continuum can be neglected.

### RADIATIVE DECAYS

$\Gamma(\gamma Xc2)/\Gamma_{\text{total}}$					$\Gamma_{98}/\Gamma$				
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<0.64	90	37 ABLIKIM	15J BES3	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma\gamma J/\psi$					
• • • We do not use the following data for averages, fits, limits, etc. • • •									
<2.0	90	38 BRIERE	06 CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma + \text{hadrons}$					
<0.9	90	39 COAN	06A CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma\gamma J/\psi$					

See key on page 601

# Meson Particle Listings

## $\psi(3770)$ , $\psi(3823)$

$\Gamma(\gamma\chi_{c1})/\Gamma_{\text{total}}$					$\Gamma_{99}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>2.48 \pm 0.23</math> OUR AVERAGE</b>					

1.9 $\pm 0.4 \pm 0.6$	202	40	ABLIKIM	16B	BES3	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma + \text{hadrons}$	
2.48 $\pm 0.15 \pm 0.23$	0.6k		ABLIKIM	15J	BES3	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma\gamma J/\psi$	
2.4 $\pm 0.8 \pm 0.2$		41	ABLIKIM	14H	BES3	$e^+e^- \rightarrow \psi(3770) \rightarrow K_S^0 K^\pm \pi^\mp$	
2.9 $\pm 0.5 \pm 0.4$		42	BRIERE	06	CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma + \text{hadrons}, \gamma\gamma J/\psi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
3.9 $\pm 1.4 \pm 0.6$	54	43	BRIERE	06	CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma + \text{hadrons}$	
2.8 $\pm 0.5 \pm 0.4$	53	39	COAN	06A	CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma\gamma J/\psi$	

$\Gamma(\gamma\chi_{c1})/\Gamma(J/\psi\pi^+\pi^-)$					$\Gamma_{99}/\Gamma_4$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.49 \pm 0.31 \pm 0.26</math></b>	$53 \pm 10$	<sup>44</sup> COAN	06A CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow$	

$\Gamma(\gamma\chi_{c0})/\Gamma_{\text{total}}$					$\Gamma_{100}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>7.0 \pm 0.6</math> OUR AVERAGE</b>					

6.9 $\pm 0.3 \pm 0.7$	2.2K	45	ABLIKIM	16B	BES3	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma + \text{hadrons}$	
7.3 $\pm 0.7 \pm 0.6$	274		BRIERE	06	CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma + \text{hadrons}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
< 44	90	39	COAN	06A	CLEO	$e^+e^- \rightarrow \psi(3770) \rightarrow \gamma\gamma J/\psi$	

$\Gamma(\gamma\chi_{c0})/\Gamma(\gamma\chi_{c2})$					$\Gamma_{100}/\Gamma_{98}$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>8	90	<sup>46</sup> BRIERE	06	CLEO	$e^+e^- \rightarrow \psi(3770)$

$\Gamma(\gamma\chi_{c0})/\Gamma(\gamma\chi_{c1})$					$\Gamma_{100}/\Gamma_{99}$
VALUE		DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$2.5 \pm 0.6$		46	BRIERE	06	CLEO $e^+e^- \rightarrow \psi(3770)$

$\Gamma(\gamma\eta_c)/\Gamma_{\text{total}}$					$\Gamma_{101}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN		
$<7 \times 10^{-4}$	90	47 ABLIKIM	14H	BES3	

$\Gamma(\gamma\eta_c(2S))/\Gamma_{\text{total}}$					$\Gamma_{102}/\Gamma$
VALUE	CL%	DOCUMENT ID		TECN	
$<9 \times 10^{-4}$	90	48	ABLIKIM	14H	BES3

$\Gamma(\gamma\eta)/\Gamma_{\text{total}}$					$\Gamma_{103}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.8</b>	90	49 PEDLAR	09 CLE3	$\psi(2S) \rightarrow \gamma X$	

$\Gamma(\gamma\eta)/\Gamma_{\text{total}}$					$\Gamma_{104}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.5</b>	90	49 PEDLAR	09 CLE3	$\psi(2S) \rightarrow \gamma X$	

$\Gamma(\gamma\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{105}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<2	90	PEDLAR	09	CLE3	$\psi(2S) \rightarrow \gamma X$

- <sup>37</sup> This limit is equivalent to  $(0.25 \pm 0.21 \pm 0.18) \times 10^{-3}$  branching fraction value.
- <sup>38</sup> Uses  $B(\psi(2S) \rightarrow \gamma\chi_{c2}) = 9.22 \pm 0.11 \pm 0.46\%$  from ATHAR 04,  $\psi(2S)$  mass and width from PDG 04, and  $\Gamma_{ee}(\psi(2S)) = 2.54 \pm 0.03 \pm 0.11$  keV from ADAM 06.
- <sup>39</sup> Using  $\Gamma_{ee}(\psi(2S)) = (2.54 \pm 0.03 \pm 0.11)$  keV from ADAM 06 and taking  $\sigma(e^+e^- \rightarrow D\bar{D})$  from HE 05 for  $\sigma(e^+e^- \rightarrow \psi(3770))$ .
- <sup>40</sup> ABLIKIM 16B reports  $(1.94 \pm 0.42 \pm 0.64) \times 10^{-3}$  from a measurement of  $[\Gamma(\psi(3770) \rightarrow \gamma\chi_{c1})/\Gamma_{\text{total}}] / [B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c1}(1P)) = (9.55 \pm 0.31) \times 10^{-2}$ .
- <sup>41</sup> ABLIKIM 14H reports  $[\Gamma(\psi(3770) \rightarrow \gamma\chi_{c1})/\Gamma_{\text{total}}] \times [B(\chi_{c1}(1P) \rightarrow K_S^0 K^\pm \pi^\mp)] = (8.51 \pm 2.39 \pm 1.42) \times 10^{-6}$  which we divide by our best value  $B(\chi_{c1}(1P) \rightarrow K_S^0 K^\pm \pi^\mp) = 0.00356 \pm 0.00030$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. We have calculated the best value of  $B(\chi_{c1}(1P) \rightarrow K_S^0 K^\pm \pi^\mp)$  as  $1/2$  of  $B(\chi_{c1}(1P) \rightarrow \bar{K}^0 K^+ \pi^- + \text{c.c.}) = (7.1 \pm 0.6) \times 10^{-3}$ .
- <sup>42</sup> Averages the two measurements from COAN 06A and BRIERE 06.
- <sup>43</sup> Uses  $B(\psi(2S) \rightarrow \gamma\chi_{c1}) = 9.07 \pm 0.11 \pm 0.54\%$  from ATHAR 04,  $\psi(2S)$  mass and width from PDG 04, and  $\Gamma_{ee}(\psi(2S)) = 2.54 \pm 0.03 \pm 0.11$  keV from ADAM 06.
- <sup>44</sup> Using  $B(\psi(3770) \rightarrow J/\psi\pi^+\pi^-) = (1.89 \pm 0.20 \pm 0.20) \times 10^{-3}$  from ADAM 06.
- <sup>45</sup> ABLIKIM 16B reports  $(6.88 \pm 0.28 \pm 0.67) \times 10^{-3}$  from a measurement of  $[\Gamma(\psi(3770) \rightarrow \gamma\chi_{c0})/\Gamma_{\text{total}}] / [B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P))]$  assuming  $B(\psi(2S) \rightarrow \gamma\chi_{c0}(1P)) = (9.99 \pm 0.27) \times 10^{-2}$ .

<sup>46</sup> Not independent of other results in BRIERE 06.

<sup>47</sup> ABLIKIM 14H reports  $[\Gamma(\psi(3770) \rightarrow \gamma\eta_c)/\Gamma_{\text{total}}] \times [B(\eta_c(1S) \rightarrow K_S^0 K^\pm \pi^\mp)] < 16 \times 10^{-6}$  which we divide by our best value  $B(\eta_c(1S) \rightarrow K_S^0 K^\pm \pi^\mp) = 2.43 \times 10^{-2}$ . We have calculated the best value of  $B(\eta_c(1S) \rightarrow K_S^0 K^\pm \pi^\mp)$  as  $1/3$  of  $B(\eta_c(1S) \rightarrow K\bar{K}\pi) = 7.3 \times 10^{-2}$ .

<sup>48</sup> ABLIKIM 14H reports  $[\Gamma(\psi(3770) \rightarrow \gamma\eta_c(2S))/\Gamma_{\text{total}}] \times [B(\eta_c(2S) \rightarrow K_S^0 K^\pm \pi^\mp)] < 5.6 \times 10^{-6}$  which we divide by our best value  $B(\eta_c(2S) \rightarrow K_S^0 K^\pm \pi^\mp) = 6 \times 10^{-3}$ . We have calculated the best value of  $B(\eta_c(2S) \rightarrow K_S^0 K^\pm \pi^\mp)$  as  $1/3$  of  $B(\eta_c(2S) \rightarrow K\bar{K}\pi) = 1.9 \times 10^{-2}$ .

<sup>49</sup> Assuming maximal destructive interference between  $\psi(3770)$  and continuum sources.

## $\psi(3770)$ REFERENCES

ABLIKIM	16B	PL B753 103	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15J	PR D91 092009	M. Ablikim <i>et al.</i>	(BES III Collab.)
DRUZHININ	15	PR D92 054024	V.P. Druzhinin	(NOVO)
ABLIKIM	14H	PR D89 112005	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	14L	PL B735 101	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	14O	PR D90 032007	M. Ablikim <i>et al.</i>	(BES III Collab.)
BONVICINI	14	PR D89 072002	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
ABLIKIM	13Q	PR D87 112011	Ablikim M. <i>et al.</i>	(BES III Collab.)
ANASHIN	12A	PL B711 292	V.V. Anashin <i>et al.</i>	(KEDR Collab.)
ABLIKIM	10D	EPJ C66 11	M. Ablikim <i>et al.</i>	(BES II Collab.)
BESSION	10	PRL 104 159901 (err.)	D. Besson <i>et al.</i>	(CLEO Collab.)
ABLIKIM	09C	EPJ C64 243	M. Ablikim <i>et al.</i>	(BES Collab.)
PEDLAR	09	PR D79 111101	T.K. Pedlar <i>et al.</i>	(CLEO Collab.)
ABLIKIM	08B	PL B659 74	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	08D	PL B660 315	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	08M	PL B670 179	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	08N	PL B670 184	M. Ablikim <i>et al.</i>	(BES Collab.)
AUBERT	08B	PR D77 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
BRODZICKA	08	PRL 100 092001	J. Brodzicka <i>et al.</i>	(BELLE Collab.)
PAKHOLOVA	08	PR D77 011103	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
ABLIKIM	07B	PL B650 111	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	07E	PL B652 238	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	07F	PL B656 30	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	07I	EPJ C52 805	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	07K	PR D76 122002	M. Ablikim <i>et al.</i>	(BES Collab.)
AUBERT	07BE	PR D76 111105	B. Aubert <i>et al.</i>	(BABAR Collab.)
DOBBS	07	PR D76 112001	S. Dobbs <i>et al.</i>	(CLEO Collab.)
ABLIKIM	06L	PRL 97 121801	M. Ablikim <i>et al.</i>	(BES Collab.)
ABLIKIM	06N	PL B641 145	M. Ablikim <i>et al.</i>	(BES Collab.)
ADAM	06	PRL 96 082004	N.E. Adam <i>et al.</i>	(CLEO Collab.)
ADAMS	06	PR D73 012002	G.S. Adams <i>et al.</i>	(CLEO Collab.)
BESSION	06	PRL 96 092002	D. Besson <i>et al.</i>	(CLEO Collab.)
Also		PRL 104 159901 (err.)	D. Besson <i>et al.</i>	(CLEO Collab.)
BRIERE	06	PR D74 031106	R.A. Briere <i>et al.</i>	(CLEO Collab.)
COAN	06A	PRL 96 182002	T.E. Coan <i>et al.</i>	(CLEO Collab.)
CRONIN-HENNESSY	06	PR D74 012005	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
HUANG	06A	PRL 96 032003	G.S. Huang <i>et al.</i>	(CLEO Collab.)
BAI	05	PL B605 63	J.Z. Bai <i>et al.</i>	(BES Collab.)
HE	05	PRL 95 121801	Q. He <i>et al.</i>	(CLEO Collab.)
Also		PRL 96 199903 (err.)	Q. He <i>et al.</i>	(CLEO Collab.)
ABLIKIM	04F	PR D70 077101	M. Ablikim <i>et al.</i>	(BES Collab.)
ATHAR	04	PR D70 112002	S.B. Athar <i>et al.</i>	(CLEO Collab.)
CHISTOV	04	PRL 93 051803	R. Chistov <i>et al.</i>	(BELLE Collab.)
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
BAI	02C	PRL 88 101802	J.Z. Bai <i>et al.</i>	(BES Collab.)
ADLER	88C	PRL 60 89	J. Adler <i>et al.</i>	(Mark III Collab.)
SCHINDLER	80	PR D21 2716	R.H. Schindler <i>et al.</i>	(Mark II Collab.)
BACINO	78	PRL 40 671	W.J. Bacino <i>et al.</i>	(SLAC, UCLA, UCI)
RAPIDIS	77	PRL 39 526	P.A. Rapidis <i>et al.</i>	(LGW Collab.)

$\psi(3823)$   
was  $X(3823)$ .

$I_G(J^{PC}) = ?^?(2^{-}-)$   
 $J, P$  need confirmation.

Seen by BHARDWAJ 13 in  $B \rightarrow \chi_{c1} \gamma K$  and ABLIKIM 15S in  $e^+e^- \rightarrow \pi^+\pi^-\gamma\chi_{c1}$  decays as a narrow peak in the invariant mass distribution of the  $\chi_{c1} \gamma$  system. Properties consistent with the  $\psi_2(1^3D_2) c\bar{c}$  state.

## $\psi(3823)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3822.2±1.2 OUR AVERAGE</b>				
3821.7±1.3±0.7	19 ± 5	<sup>1</sup> ABLIKIM	15s BES3	$e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}\gamma$
3823.1±1.8±0.7	33 ± 10	<sup>2</sup> BHARDWAJ	13 BELL	$B \rightarrow \chi_{c1}\gamma K$

- <sup>1</sup> From a simultaneous unbinned maximum likelihood fit of  $e^+e^- \rightarrow \pi^+\pi^-\chi_{c1} \gamma$  data (the  $\pi^+\pi^-$  recoil mass) taken at  $\sqrt{s}$  values of 4.23, 4.26, 4.36, 4.42, and 4.60 GeV to simulated events including both  $\psi(2S) \rightarrow \chi_{c1} \gamma$  and  $\psi(3823) \rightarrow \chi_{c1} \gamma$  together, with floating mass scale offset for  $\psi(2S)$ , floating  $\psi(3823)$  mass, and zero  $\psi(3823)$  width, resulting in a significance of 5.9 $\sigma$  when including systematic uncertainties.
- <sup>2</sup> From a simultaneous fit to  $B^\pm \rightarrow (\chi_{c1} \gamma) K^\pm$  and  $B^0 \rightarrow (\chi_{c1} \gamma) K_S^0$  with significance 4.0 $\sigma$  including systematics. Corrected for the measured  $\psi(2S)$  mass using  $B \rightarrow \psi(2S) K \rightarrow (\gamma\chi_{c1}) K$  decays.

## $\psi(3823)$ WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT	
<16	90	<sup>1</sup> ABLIKIM	15s	BES3	$e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<24	90	<sup>2</sup> BHARDWAJ	13	BELL	$B \rightarrow \chi_{c1}\gamma K$

# Meson Particle Listings

## $\psi(3823)$ , $X(3872)$

<sup>1</sup> From a fit of  $e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}\gamma$  data (the  $\pi^+\pi^-$  recoil mass) taken at  $\sqrt{s}$  values of 4.23, 4.26, 4.36, 4.42, and 4.60 GeV to a Breit-Wigner function with the mass fixed from the likelihood fit above, Gaussian resolution smearing, and floating width.

<sup>2</sup> From a simultaneous fit to  $B^\pm \rightarrow (\chi_{c1}\gamma)K^\pm$  and  $B^0 \rightarrow (\chi_{c1}\gamma)K_S^0$  with significance 4.0 $\sigma$  including systematics.

### $\psi(3823)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\chi_{c1}\gamma$	seen
$\Gamma_2$ $\chi_{c2}\gamma$	not seen

### $\psi(3823)$ BRANCHING RATIOS

$\Gamma(\chi_{c1}\gamma)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	$33 \pm 10$	<sup>1</sup> BHARDWAJ	13	BELL $B^+ \rightarrow \chi_{c1}\gamma K^+$

<sup>1</sup> Reported  $B(B^\pm \rightarrow \psi(3823)K^\pm) \times B(\psi(3823) \rightarrow \gamma\chi_{c1}) = (9.7 \pm 2.8 \pm 1.1) \times 10^{-6}$  with statistical significance 3.8 $\sigma$ .

$\Gamma(\chi_{c2}\gamma)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
not seen	<sup>1</sup> ABLIKIM	15s	BES3 $e^+e^- \rightarrow \pi^+\pi^-\chi_{c2}\gamma$
not seen	<sup>2</sup> BHARDWAJ	13	BELL $B^+ \rightarrow \chi_{c2}\gamma K^+$

<sup>1</sup> From a simultaneous unbinned maximum likelihood fit of  $e^+e^- \rightarrow \pi^+\pi^-\chi_{c2}\gamma$  data (the  $\pi^+\pi^-$  recoil mass) taken at  $\sqrt{s}$  values of 4.23, 4.26, 4.36, 4.42, and 4.60 GeV to simulated events including both  $\psi(2S) \rightarrow \chi_{c2}\gamma$  and  $\psi(3823) \rightarrow \chi_{c2}\gamma$  together, with floating mass scale offset for  $\psi(2S)$ ,  $\psi(3823)$  mass floating (fixed to that above), and zero  $\psi(3823)$  width.

<sup>2</sup> Reported  $B(B^\pm \rightarrow \psi(3823)K^\pm) \times B(\psi(3823) \rightarrow \gamma\chi_{c2}) < 3.6 \times 10^{-6}$  at 90% CL.

$\Gamma(\chi_{c2}\gamma)/\Gamma(\chi_{c1}\gamma)$				$\Gamma_2/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.41	90	BHARDWAJ 13	BELL	$B^+ \rightarrow \chi_{c1/c2}\gamma K^+$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.42	90	<sup>1</sup> ABLIKIM	15s	BES3 $e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}\gamma$
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<sup>1</sup> From a simultaneous unbinned maximum likelihood fit of  $e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}(2)\gamma$  data (the  $\pi^+\pi^-$  recoil mass) taken at  $\sqrt{s}$  values of 4.23, 4.26, 4.36, 4.42, and 4.60 GeV to simulated events including both  $\psi(2S) \rightarrow \chi_{c1}(2)\gamma$  and  $\psi(3823) \rightarrow \chi_{c1}(2)\gamma$  together, with floating mass scale offset for  $\psi(2S)$ ,  $\psi(3823)$  mass floating (fixed to that above), and zero  $\psi(3823)$  width.

### $\psi(3823)$ REFERENCES

ABLIKIM	15S	PRL 115 011803	M. Ablikim <i>et al.</i>	(BES III Collab.)
BHARDWAJ	13	PRL 111 032001	V. Bhardwaj <i>et al.</i>	(BELLE Collab.)

## X(3872)

$$I^G(J^{PC}) = 0^+(1^{++})$$

First observed by CHOI 03 in  $B \rightarrow K\pi^+\pi^-J/\psi(1S)$  decays as a narrow peak in the invariant mass distribution of the  $\pi^+\pi^-J/\psi(1S)$  final state. Isovector hypothesis excluded by AUBERT 05B and CHOI 11.

AAIJ 13Q perform a full five-dimensional amplitude analysis of the angular correlations between the decay products in  $B^+ \rightarrow X(3872)K^+$  decays, where  $X(3872) \rightarrow J/\psi\pi^+\pi^-$  and  $J/\psi \rightarrow \mu^+\mu^-$ , which unambiguously gives the  $J^{PC} = 1^{++}$  assignment under the assumption that the  $\pi^+\pi^-$  and  $J/\psi$  are in an  $S$ -wave. AAIJ 15AO extend this analysis with more data to limit  $D$ -wave contributions to < 4% at 95% CL.

See our note on “Developments in Heavy Quarkonium Spectroscopy”.

### X(3872) MASS FROM $J/\psi X$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3871.69 <math>\pm</math> 0.17 OUR AVERAGE</b>				
3871.9 $\pm$ 0.7 $\pm$ 0.2	20 $\pm$ 5	ABLIKIM	14	BES3 $e^+e^- \rightarrow J/\psi\pi^+\pi^-\gamma$
3871.95 $\pm$ 0.48 $\pm$ 0.12	0.6k	AAIJ	12H	LHCB $pp \rightarrow J/\psi\pi^+\pi^-X$
3871.85 $\pm$ 0.27 $\pm$ 0.19	$\sim$ 170	<sup>1</sup> CHOI	11	BELL $B \rightarrow K\pi^+\pi^-J/\psi$
3873 $\pm$ 1.8 $\pm$ 1.3	27 $\pm$ 8	<sup>2</sup> DEL-AMO-SA.10B	BABR	$B \rightarrow \omega J/\psi K$
3871.61 $\pm$ 0.16 $\pm$ 0.19	6k	<sup>2,3</sup> AALTONEN	09AU	CDF2 $p\bar{p} \rightarrow J/\psi\pi^+\pi^-X$
3871.4 $\pm$ 0.6 $\pm$ 0.1	93.4	AUBERT	08Y	BABR $B^+ \rightarrow K^+J/\psi\pi^+\pi^-$
3868.7 $\pm$ 1.5 $\pm$ 0.4	9.4	AUBERT	08Y	BABR $B^0 \rightarrow K_S^0J/\psi\pi^+\pi^-$
3871.8 $\pm$ 3.1 $\pm$ 3.0	522	<sup>2,4</sup> ABAZOV	04F	D0 $p\bar{p} \rightarrow J/\psi\pi^+\pi^-X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

3868.6 $\pm$ 1.2 $\pm$ 0.2	8	<sup>5</sup> AUBERT	06	BABR $B^0 \rightarrow K_S^0J/\psi\pi^+\pi^-$
3871.3 $\pm$ 0.6 $\pm$ 0.1	61	<sup>5</sup> AUBERT	06	BABR $B^- \rightarrow K^-J/\psi\pi^+\pi^-$
3873.4 $\pm$ 1.4	25	<sup>6</sup> AUBERT	05R	BABR $B^+ \rightarrow K^+J/\psi\pi^+\pi^-$
3871.3 $\pm$ 0.7 $\pm$ 0.4	730	<sup>2,7</sup> ACOSTA	04	CDF2 $p\bar{p} \rightarrow J/\psi\pi^+\pi^-X$
3872.0 $\pm$ 0.6 $\pm$ 0.5	36	<sup>8</sup> CHOI	03	BELL $B \rightarrow K\pi^+\pi^-J/\psi$
3836 $\pm$ 13	58	<sup>2,9</sup> ANTONIAZZI	94	E705 $300\pi^\pm\text{Li} \rightarrow J/\psi\pi^+\pi^-X$

<sup>1</sup> The mass difference for the  $X(3872)$  produced in  $B^+$  and  $B^0$  decays is  $(-0.71 \pm 0.96 \pm 0.19)$  MeV.

<sup>2</sup> Width consistent with detector resolution.

<sup>3</sup> A possible equal mixture of two states with a mass difference greater than 3.6 MeV/ $c^2$  is excluded at 95% CL.

<sup>4</sup> Calculated from the corresponding  $m_{X(3872)} - m_{J/\psi}$  using  $m_{J/\psi} = 3096.916$  MeV.

<sup>5</sup> Calculated from the corresponding  $m_{X(3872)} - m_{\psi(2S)}$  using  $m_{\psi(2S)} = 3686.093$  MeV. Superseded by AUBERT 08Y.

<sup>6</sup> Calculated from the corresponding  $m_{X(3872)} - m_{\psi(2S)}$  using  $m_{\psi(2S)} = 3685.96$  MeV. Superseded by AUBERT 06.

<sup>7</sup> Superseded by AALTONEN 09AU.

<sup>8</sup> Superseded by CHOI 11.

<sup>9</sup> A lower mass value can be due to an incorrect momentum scale for soft pions.

### X(3872) MASS FROM $\bar{D}^{*0}D^0$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3872.9 $\pm$ 0.6 $\pm$ 0.4 $\pm$ 0.4 $\pm$ 0.5	50	<sup>1,2</sup> AUSHEV	10	BELL $B \rightarrow \bar{D}^{*0}D^0K$
3875.1 $\pm$ 0.7 $\pm$ 0.5	33 $\pm$ 6	<sup>2</sup> AUBERT	08B	BABR $B \rightarrow \bar{D}^{*0}D^0K$
3875.2 $\pm$ 0.7 $\pm$ 0.9 $\pm$ 1.8	24 $\pm$ 6	<sup>2,3</sup> GOKHROO	06	BELL $B \rightarrow D^0\bar{D}^0\pi^0K$

<sup>1</sup> Calculated from the measured  $m_{X(3872)} - m_{D^{*0}} - m_{\bar{D}^0} = 1.1^{+0.6+0.1}_{-0.4-0.3}$  MeV.

<sup>2</sup> Experiments report  $D^{*0}\bar{D}^0$  invariant mass above  $D^{*0}\bar{D}^0$  threshold because  $D^{*0}$  decay products are kinematically constrained to the  $D^{*0}$  mass, even though the  $D^{*0}$  may decay off-shell.

<sup>3</sup> Superseded by AUSHEV 10.

### $m_{X(3872)} - m_{J/\psi}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>774.9 <math>\pm</math> 3.1 <math>\pm</math> 3.0</b>	522	ABAZOV	04F	D0 $p\bar{p} \rightarrow J/\psi\pi^+\pi^-X$

### $m_{X(3872)} - m_{\psi(2S)}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
187.4 $\pm$ 1.4	25	<sup>1</sup> AUBERT	05R	BABR $B^+ \rightarrow K^+J/\psi\pi^+\pi^-$

<sup>1</sup> Superseded by AUBERT 06.

### X(3872) WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<1.2	90		CHOI	11	BELL $B \rightarrow K\pi^+\pi^-J/\psi$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<2.4	90		ABLIKIM	14	BES3 $e^+e^- \rightarrow J/\psi\pi^+\pi^-\gamma$
<3.3	90		AUBERT	08Y	BABR $B^+ \rightarrow K^+J/\psi\pi^+\pi^-$
<4.1	90	69	AUBERT	06	BABR $B \rightarrow K\pi^+\pi^-J/\psi$
<2.3	90	36	<sup>1</sup> CHOI	03	BELL $B \rightarrow K\pi^+\pi^-J/\psi$

<sup>1</sup> Superseded by CHOI 11.

### X(3872) WIDTH FROM $\bar{D}^{*0}D^0$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.9 $\pm$ 2.8 $\pm$ 0.2 $\pm$ 1.4 $\pm$ 1.1	50	<sup>1</sup> AUSHEV	10	BELL $B \rightarrow \bar{D}^{*0}D^0K$
3.0 $\pm$ 1.9 $\pm$ 0.9 $\pm$ 1.4	33 $\pm$ 6	AUBERT	08B	BABR $B \rightarrow \bar{D}^{*0}D^0K$

<sup>1</sup> With a measured value of  $B(B \rightarrow X(3872)K) \times B(X(3872) \rightarrow D^{*0}\bar{D}^0) = (0.80 \pm 0.20 \pm 0.10) \times 10^{-4}$ , assumed to be equal for both charged and neutral modes.

### X(3872) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $e^+e^-$	
$\Gamma_2$ $\pi^+\pi^-J/\psi(1S)$	> 2.6 %
$\Gamma_3$ $\rho^0J/\psi(1S)$	
$\Gamma_4$ $\omega J/\psi(1S)$	> 1.9 %
$\Gamma_5$ $D^0\bar{D}^0\pi^0$	>32 %
$\Gamma_6$ $\bar{D}^{*0}D^0$	>24 %
$\Gamma_7$ $\gamma\gamma$	



$\Gamma_8$	$D^0 \bar{D}^0$	
$\Gamma_9$	$D^+ D^-$	
$\Gamma_{10}$	$\gamma \chi_{c1}$	
$\Gamma_{11}$	$\gamma \chi_{c2}$	
$\Gamma_{12}$	$\gamma J/\psi$	$> 6 \times 10^{-3}$
$\Gamma_{13}$	$\gamma \psi(2S)$	$> 3.0 \%$
$\Gamma_{14}$	$\pi^+ \pi^- \eta_c(1S)$	not seen
$\Gamma_{15}$	$\rho \bar{\rho}$	not seen
<b>C-violating decays</b>		
$\Gamma_{16}$	$\eta J/\psi$	

**X(3872) PARTIAL WIDTHS**

$\Gamma(e^+ e^-)$					$\Gamma_1$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 4.3$	90	<sup>1</sup> ABLIKIM	15v BES3	$4.0\text{--}4.4 e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
$< 280$	90	<sup>2</sup> YUAN	04 RVUE	$e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
<sup>1</sup> ABLIKIM 15v reports this limit from the measurement of $\Gamma(X(3872) \rightarrow \pi^+ \pi^- J/\psi(1S)) \times \Gamma(X(3872) \rightarrow e^+ e^-)/\Gamma < 0.13$ eV using $\Gamma(X(3872) \rightarrow \pi^+ \pi^- J/\psi(1S))/\Gamma = 3\%$ .					
<sup>2</sup> Using BAI 98E data on $e^+ e^- \rightarrow \pi^+ \pi^- \ell^+ \ell^-$ . Assuming that $\Gamma(\pi^+ \pi^- J/\psi)$ of X(3872) is the same as that of $\psi(2S)$ (85.4 keV).					

**X(3872)  $\Gamma(i)\Gamma(e^+ e^-)/\Gamma(\text{total})$** 

$\Gamma(\pi^+ \pi^- J/\psi(1S)) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_2 \Gamma_1/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 0.13</b>	90	ABLIKIM	15v BES3	$4.0\text{--}4.4 e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 6.2$	90	<sup>1,2</sup> AUBERT	05D BABR	$10.6 e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
$< 8.3$	90	<sup>2</sup> DOBBS	05 CLE3	$e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
$< 10$	90	<sup>3</sup> YUAN	04 RVUE	$e^+ e^- \rightarrow \pi^+ \pi^- J/\psi$	
<sup>1</sup> Using $B(X(3872) \rightarrow J/\psi \pi^+ \pi^-) \cdot B(J/\psi \rightarrow \mu^+ \mu^-) \cdot \Gamma(X(3872) \rightarrow e^+ e^-) < 0.37$ eV from AUBERT 05D and $B(J/\psi \rightarrow \mu^+ \mu^-) = 0.0588 \pm 0.0010$ from the PDG 04.					
<sup>2</sup> Assuming X(3872) has $J^{PC} = 1^{--}$ .					
<sup>3</sup> Using BAI 98E data on $e^+ e^- \rightarrow \pi^+ \pi^- \ell^+ \ell^-$ . From theoretical calculation of the production cross section and using $B(J/\psi \rightarrow \mu^+ \mu^-) = (5.88 \pm 0.10)\%$ .					

**X(3872)  $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$** 

$\Gamma(\pi^+ \pi^- J/\psi(1S)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_2 \Gamma_7/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 12.9$	90	<sup>1</sup> DOBBS	05 CLE3	$e^+ e^- \rightarrow \pi^+ \pi^- J/\psi \gamma$	
<sup>1</sup> Assuming X(3872) has positive C parity and spin 0.					

$\Gamma(\omega J/\psi(1S)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_4 \Gamma_7/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 1.7$	90	<sup>1</sup> LEES	12AD BABR	$e^+ e^- \rightarrow e^+ e^- \omega J/\psi$	
<sup>1</sup> Assuming X(3872) has spin 2.					

$\Gamma(\pi^+ \pi^- \eta_c(1S)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_{14} \Gamma_7/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 11.1</b>	90	LEES	12AE BABR	$e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \eta_c$	

**X(3872) BRANCHING RATIOS**

$\Gamma(\pi^+ \pi^- J/\psi(1S))/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&gt; 0.026</b>	$93 \pm 17$	<sup>1</sup> AUBERT	08y BABR	$B \rightarrow X(3872) K$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen	151	<sup>2</sup> BALA	15 BELL	$B \rightarrow X(3872) K \pi$	
$> 0.04$	30	<sup>3</sup> AUBERT	05R BABR	$B^+ \rightarrow K^+ \pi^+ \pi^- J/\psi$	
$> 0.04$	$36 \pm 7$	<sup>4</sup> CHOI	03 BELL	$B^+ \rightarrow K^+ \pi^+ \pi^- J/\psi$	
<sup>1</sup> AUBERT 08y reports $[\Gamma(X(3872) \rightarrow \pi^+ \pi^- J/\psi(1S))/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (8.4 \pm 1.5 \pm 0.7) \times 10^{-6}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					
<sup>2</sup> BALA 15 reports $B(X(3872) \rightarrow \pi^+ \pi^- J/\psi) \times B(B^0 \rightarrow X(3872) K^+ \pi^-) = (7.9 \pm 1.3 \pm 0.4) \times 10^{-6}$ and $B(X(3872) \rightarrow \pi^+ \pi^- J/\psi) \times B(B^+ \rightarrow X(3872) K^0 \pi^+) = (10.6 \pm 3.0 \pm 0.9) \times 10^{-6}$ .					
<sup>3</sup> Superseded by AUBERT 08y. AUBERT 05R reports $[\Gamma(X(3872) \rightarrow \pi^+ \pi^- J/\psi(1S))/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (1.28 \pm 0.41) \times 10^{-5}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					
<sup>4</sup> CHOI 03 reports $[\Gamma(X(3872) \rightarrow \pi^+ \pi^- J/\psi(1S))/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) / [B(B^+ \rightarrow \psi(2S) K^+) / [B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-)]] = 0.063 \pm 0.012 \pm 0.007$ which we multiply or divide by our best values $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ , $B(B^+ \rightarrow \psi(2S) K^+) = (6.26 \pm 0.24) \times 10^{-4}$ , $B(\psi(2S) \rightarrow J/\psi(1S) \pi^+ \pi^-) = (34.49 \pm 0.30) \times 10^{-2}$ .					

$\Gamma(\omega J/\psi(1S))/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&gt; 0.019</b>	$21 \pm 7$	<sup>1</sup> DEL-AMO-SA..10B	BABR	$B^+ \rightarrow \omega J/\psi K^+$	

<sup>1</sup> DEL-AMO-SANCHEZ 10B reports  $[\Gamma(X(3872) \rightarrow \omega J/\psi(1S))/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (6 \pm 2 \pm 1) \times 10^{-6}]$  which we divide by our best value  $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ . DEL-AMO-SANCHEZ 10B also reports  $B(B^0 \rightarrow X(3872) K^0) \times B(X(3872) \rightarrow J/\psi \omega) = (6 \pm 3 \pm 1) \times 10^{-6}$ .

$\Gamma(\omega J/\psi(1S))/\Gamma(\pi^+ \pi^- J/\psi(1S))$					$\Gamma_4/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.8 ± 0.3</b>	<sup>1</sup> DEL-AMO-SA..10B	BABR	$B \rightarrow \omega J/\psi K$		

<sup>1</sup> Statistical and systematic errors added in quadrature. Uses the values of  $B(B \rightarrow X(3872) K) \times B(X(3872) \rightarrow J/\psi \pi^+ \pi^-)$  reported in AUBERT 08y, taking into account the common systematics.

$\Gamma(D^0 \bar{D}^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&gt; 0.32</b>	$17 \pm 5$	<sup>1</sup> GOKHROO	06 BELL	$B^+ \rightarrow D^0 \bar{D}^0 \pi^0 K^+$	
<sup>1</sup> GOKHROO 06 reports $[\Gamma(X(3872) \rightarrow D^0 \bar{D}^0 \pi^0)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (1.02 \pm 0.31 \pm 0.21 \pm 0.29) \times 10^{-4}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					

$\Gamma(\bar{D}^{*0} D^0)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&gt; 0.24</b>	$41 \pm 9$	<sup>1</sup> AUSHEV	10 BELL	$B^+ \rightarrow D^{*0} \bar{D}^0 K^+$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$> 0.5$	$27 \pm 6$	<sup>2</sup> AUBERT	08B BABR	$B^+ \rightarrow \bar{D}^{*0} D^0 K^+$	
<sup>1</sup> AUSHEV 10 reports $[\Gamma(X(3872) \rightarrow \bar{D}^{*0} D^0)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (0.77 \pm 0.16 \pm 0.10) \times 10^{-4}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					
<sup>2</sup> AUBERT 08B reports $[\Gamma(X(3872) \rightarrow \bar{D}^{*0} D^0)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (1.67 \pm 0.36 \pm 0.47) \times 10^{-4}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					

$\Gamma(D^0 \bar{D}^0 \pi^0)/\Gamma(\pi^+ \pi^- J/\psi(1S))$					$\Gamma_5/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	<sup>1</sup> GOKHROO	06 BELL	$B \rightarrow D^0 \bar{D}^0 \pi^0 K$		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen	AUSHEV	10 BELL	$B \rightarrow D^0 \bar{D}^0 \pi^0 K$		
<sup>1</sup> May not necessarily be the same state as that observed in the $J/\psi \pi^+ \pi^-$ mode. Supersedes CHISTOV 04.					

$\Gamma(D^0 \bar{D}^0)/\Gamma(\pi^+ \pi^- J/\psi(1S))$					$\Gamma_8/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
not seen	CHISTOV	04 BELL	$B \rightarrow K D^0 \bar{D}^0$		

$\Gamma(D^+ D^-)/\Gamma(\pi^+ \pi^- J/\psi(1S))$					$\Gamma_9/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
not seen	CHISTOV	04 BELL	$B \rightarrow K D^+ D^-$		

$\Gamma(\gamma \chi_{c1})/\Gamma(\pi^+ \pi^- J/\psi(1S))$					$\Gamma_{10}/\Gamma_2$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
not seen		<sup>1</sup> BHARDWAJ	13 BELL	$B^+ \rightarrow \chi_{c1} \gamma K^+$	
<b>&lt; 0.89</b>	90	CHOI	03 BELL	$B \rightarrow K \pi^+ \pi^- J/\psi$	
<sup>1</sup> Reported $B(B^\pm \rightarrow X(3872) K^\pm) \times B(X(3872) \rightarrow \gamma \chi_{c1}) < 1.9 \times 10^{-6}$ at 90% CL.					

$\Gamma(\gamma \chi_{c2})/\Gamma(\pi^+ \pi^- J/\psi(1S))$					$\Gamma_{11}/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT		
not seen	<sup>1</sup> BHARDWAJ	13 BELL	$B^\pm \rightarrow \chi_{c2} \gamma K^\pm$		
<sup>1</sup> Reported $B(B^\pm \rightarrow X(3872) K^\pm) \times B(X(3872) \rightarrow \gamma \chi_{c2}) < 6.7 \times 10^{-6}$ at 90% CL.					

$\Gamma(\gamma J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&gt; 6 <math>\times 10^{-3}</math></b>		<sup>1</sup> BHARDWAJ	11 BELL	$B^\pm \rightarrow \gamma J/\psi K^\pm$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$> 9 \times 10^{-3}$	20	<sup>2</sup> AUBERT	09B BABR	$B^+ \rightarrow \gamma J/\psi K^+$	
$> 0.010$	19	<sup>3</sup> AUBERT, BE	06M BABR	$B^+ \rightarrow \gamma J/\psi K^+$	
<sup>1</sup> BHARDWAJ 11 reports $[\Gamma(X(3872) \rightarrow \gamma J/\psi)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (1.78 \pm 0.48 \pm 0.12) \times 10^{-6}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					
<sup>2</sup> AUBERT 09B reports $[\Gamma(X(3872) \rightarrow \gamma J/\psi)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (2.8 \pm 0.8 \pm 0.1) \times 10^{-6}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					
<sup>3</sup> Superseded by AUBERT 09B. AUBERT, BE 06M reports $[\Gamma(X(3872) \rightarrow \gamma J/\psi)/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+) = (3.3 \pm 1.0 \pm 0.3) \times 10^{-6}]$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .					

# Meson Particle Listings

## X(3872), X(3900)

$\Gamma(\gamma\psi(2S))/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
seen	36 ± 9	<sup>1</sup> AAIJ	14AH LHCb	$B^+ \rightarrow \gamma\psi(2S) K^+$	
<b>&gt;0.030</b>	25 ± 7	<sup>2</sup> AUBERT	09B BABR	$B^+ \rightarrow \gamma\psi(2S) K^+$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

not seen <sup>3</sup> BHARDWAJ 11 BELL  $B^+ \rightarrow \gamma\psi(2S) K^+$

<sup>1</sup> From  $36.4 \pm 9.0$  events of  $X(3872) \rightarrow J/\psi\gamma$  decays with a statistical significance of  $4.4\sigma$ .

<sup>2</sup> AUBERT 09B reports  $[\Gamma(X(3872) \rightarrow \gamma\psi(2S))/\Gamma_{\text{total}}] \times [B(B^+ \rightarrow X(3872) K^+)] = (9.5 \pm 2.7 \pm 0.6) \times 10^{-6}$  which we divide by our best value  $B(B^+ \rightarrow X(3872) K^+) < 3.2 \times 10^{-4}$ .

<sup>3</sup> BHARDWAJ 11 reports  $B(B^+ \rightarrow K^+ X(3872)) \times B(X \rightarrow \gamma\psi(2S)) < 3.45 \times 10^{-6}$  at 90% CL.

$\Gamma(\gamma\psi(2S))/\Gamma(\gamma J/\psi)$					$\Gamma_{13}/\Gamma_{12}$
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.6 ± 0.6 OUR AVERAGE</b>					
2.46 ± 0.64 ± 0.29	36 ± 9	<sup>1</sup> AAIJ	14AH LHCb	$B^+ \rightarrow \gamma\psi(2S) K^+$	
3.4 ± 1.4		AUBERT	09B BABR	$B^+ \rightarrow \gamma\psi(2S) K^+$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<2.1 90 BHARDWAJ 11 BELL  $B^+ \rightarrow \gamma\psi(2S) K^+$

<sup>1</sup> From  $36.4 \pm 9.0$  events of  $X(3872) \rightarrow J/\psi\gamma$  decays with a statistical significance of  $4.4\sigma$ .

$\Gamma(p\overline{p})/\Gamma(\pi^+\pi^- J/\psi(1S))$					$\Gamma_{15}/\Gamma_2$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.0 × 10<sup>-3</sup></b>	95	<sup>1</sup> AAIJ	13s LHCb	$B^+ \rightarrow p\overline{p} K^+$	
<sup>1</sup> AAIJ 13s reports $[\Gamma(X(3872) \rightarrow p\overline{p})/\Gamma(X(3872) \rightarrow \pi^+\pi^- J/\psi(1S))] \times [B(B^+ \rightarrow X(3872) K^+, X \rightarrow J/\psi\pi^+\pi^-)] < 1.7 \times 10^{-8}$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+, X \rightarrow J/\psi\pi^+\pi^-) = 8.6 \times 10^{-6}$ .					

### C-violating decays

$\Gamma(\eta J/\psi)/\Gamma(\pi^+\pi^- J/\psi(1S))$					$\Gamma_{16}/\Gamma_2$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.4</b>	90	<sup>1,2</sup> IWASHITA	14 BELL	$B \rightarrow K\eta J/\psi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.6 90 AUBERT 04Y BABR				$B \rightarrow K\eta J/\psi$	
<sup>1</sup> IWASHITA 14 reports $[\Gamma(X(3872) \rightarrow \eta J/\psi)/\Gamma(X(3872) \rightarrow \pi^+\pi^- J/\psi(1S))] \times [B(B^+ \rightarrow X(3872) K^+, X \rightarrow J/\psi\pi^+\pi^-)] < 3.8 \times 10^{-6}$ which we divide by our best value $B(B^+ \rightarrow X(3872) K^+, X \rightarrow J/\psi\pi^+\pi^-) = 8.6 \times 10^{-6}$ .					
<sup>2</sup> IWASHITA 14 also scans the $\eta J/\psi$ mass range 3.8–4.75 GeV and sets upper limits for $B(B^\pm \rightarrow X(3872) K^\pm) \times B(X(3872) \rightarrow \eta J/\psi)$ in 5 MeV intervals.					

### X(3872) REFERENCES

AAIJ	15AO	PR D92 011102	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABLIKIM	15V	PL B749 414	M. Ablikim <i>et al.</i>	(BES III Collab.)
BALA	15	PR D91 051101	A. Bala <i>et al.</i>	(Belle Collab.)
AAIJ	14AH	NP B886 665	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABLIKIM	14	PRL 112 092001	M. Ablikim <i>et al.</i>	(BES III Collab.)
IWASHITA	14	PTEP 2014 043C01	T. Iwashita <i>et al.</i>	(Belle Collab.)
AAIJ	13Q	PRL 110 222001	R. Aaij <i>et al.</i>	(LHCb Collab.) JP
AAIJ	13S	EPJ C73 2462	R. Aaij <i>et al.</i>	(LHCb Collab.)
BHARDWAJ	13	PRL 111 032001	V. Bhardwaj <i>et al.</i>	(Belle Collab.)
AAIJ	12H	EPJ C72 1972	R. Aaij <i>et al.</i>	(LHCb Collab.)
LEES	12AD	PR D86 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	12AE	PR D86 092005	J.P. Lees <i>et al.</i>	(BABAR Collab.)
BHARDWAJ	11	PRL 107 091903	V. Bhardwaj <i>et al.</i>	(Belle Collab.)
CHOI	11	PR D94 052004	S.-K. Choi <i>et al.</i>	(Belle Collab.)
AUSHEV	10	PR D81 031103	T. Aushev <i>et al.</i>	(Belle Collab.)
DEL-AMO-SA...	10B	PR D82 011101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
AALTONEN	09AU	PRL 103 152001	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AUBERT	09B	PRL 102 132001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08B	PR D77 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08Y	PR D77 111101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	06	PR D73 011101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,BE	06M	PR D74 071101	B. Aubert <i>et al.</i>	(BABAR Collab.)
GOKHROO	06	PRL 97 162002	G. Gokhroo <i>et al.</i>	(Belle Collab.)
AUBERT	05B	PR D71 031501	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	05D	PR D71 052001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	05R	PR D71 071103	B. Aubert <i>et al.</i>	(BABAR Collab.)
DOBBS	05	PRL 94 032004	S. Dobbs <i>et al.</i>	(CLEO Collab.)
ABAZOV	04F	PRL 93 162002	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ACOSTA	04	PRL 93 072001	D. Acosta <i>et al.</i>	(CDF Collab.)
AUBERT	04Y	PRL 93 041801	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHISTOV	04	PRL 93 051803	R. Chistov <i>et al.</i>	(Belle Collab.)
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	(PDG Collab.)
YUAN	04	PL B579 74	C.Z. Yuan <i>et al.</i>	
CHOI	03	PRL 91 262001	S.-K. Choi <i>et al.</i>	(Belle Collab.)
BAI	98E	PR D57 3854	J.Z. Bai <i>et al.</i>	(BES Collab.)
ANTONIAZZI	94	PR D50 4258	L. Antoniazzi <i>et al.</i>	(E705 Collab.)

## X(3900)

$I^G(J^{PC}) = 1^+(1^{+-})$

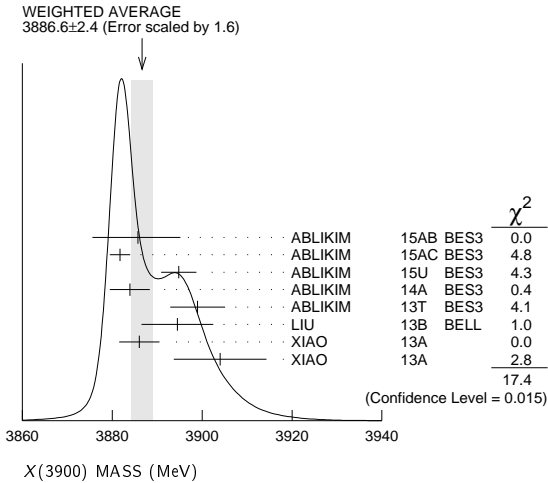
Charged X(3900) seen as a peak in the invariant mass distribution of the  $J/\psi\pi^\pm$  system by BES III (ABLIKIM 13T) in  $e^+e^- \rightarrow \pi^+\pi^- J/\psi$  at c.m. energy of 4.26 GeV and by radiative return from  $e^+e^-$  collisions at  $\sqrt{s}$  from 9.46 to 10.86 GeV at Belle (LIU 13B). Angular analysis of ABLIKIM 14A and ABLIKIM 15AC favor the  $J^P = 1^+$  assignment. Neutral X(3900) seen in the  $J/\psi\pi^0$  invariant mass distribution in  $e^+e^- \rightarrow \pi^0\pi^0 J/\psi$  at c.m. energies of 4.23, 4.26, and 4.36 GeV by BES III (ABLIKIM 15U) and at 4.17 GeV by XIAO 13A. Peaks in  $(D\overline{D}^*)^{0,\pm}$  reported by BES III (ABLIKIM 14A, ABLIKIM 15AB) are assumed to be related.

### X(3900) MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>3886.6 ± 2.4 OUR AVERAGE</b>					Error includes scale factor of 1.6. See the ideogram below.
3885.7 <sup>+4.3</sup> <sub>-5.7</sub> ± 8.4		<sup>1</sup> ABLIKIM	15AB BES3	0	$e^+e^- \rightarrow \pi^0(D\overline{D}^*)^0$
3881.7 ± 1.6 ± 1.6	1248	<sup>1</sup> ABLIKIM	15AC BES3	±	$e^+e^- \rightarrow \pi^\pm(D\overline{D}^*)^\mp$
3894.8 ± 2.3 ± 3.2	356	<sup>1</sup> ABLIKIM	15U BES3	0	$e^+e^- \rightarrow \pi^0\pi^0 J/\psi$
3883.9 ± 1.5 ± 4.2	1212	<sup>1</sup> ABLIKIM	14A BES3	±	$e^+e^- \rightarrow \pi^\pm(D\overline{D}^*)^\mp$
3899.0 ± 3.6 ± 4.9	307	<sup>1</sup> ABLIKIM	13T BES3	±	$e^+e^- \rightarrow \pi^+\pi^- J/\psi$
3894.5 ± 6.6 ± 4.5	159	<sup>1</sup> LIU	13B BELL	±	$e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$
3886 ± 4 ± 2	81	<sup>1,2</sup> XIAO	13A	±	4.17 $e^+e^- \rightarrow \pi^+\pi^- J/\psi$
3904 ± 9 ± 5	25	<sup>1,2</sup> XIAO	13A	0	4.17 $e^+e^- \rightarrow \pi^0\pi^0 J/\psi$

<sup>1</sup> Neglecting interference between the X(3900) and non-resonant continuum.

<sup>2</sup> For  $M^2(\pi^+\pi^-) < 0.65 \text{ GeV}^2$ . Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration.



### X(3900) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>28.1 ± 2.1 OUR AVERAGE</b>					
35 <sup>+11</sup> <sub>-12</sub> ± 15		<sup>1</sup> ABLIKIM	15AB BES3	0	$e^+e^- \rightarrow \pi^0(D\overline{D}^*)^0$
26.6 ± 2.0 ± 2.1	1248	<sup>1</sup> ABLIKIM	15AC BES3	±	$e^+e^- \rightarrow \pi^\pm(D\overline{D}^*)^\mp$
29.6 ± 8.2 ± 8.2	356	<sup>1</sup> ABLIKIM	15U BES3	0	$e^+e^- \rightarrow \pi^0\pi^0 J/\psi$
24.8 ± 3.3 ± 11.0	1212	<sup>1</sup> ABLIKIM	14A BES3	±	$e^+e^- \rightarrow \pi^\pm(D\overline{D}^*)^\mp$
46 ± 10 ± 20	307	<sup>1</sup> ABLIKIM	13T BES3	±	$e^+e^- \rightarrow \pi^+\pi^- J/\psi$
63 ± 24 ± 26	159	<sup>1</sup> LIU	13B BELL	±	$e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$
37 ± 4 ± 8	81	<sup>1,2</sup> XIAO	13A	±	4.17 $e^+e^- \rightarrow \pi^+\pi^- J/\psi$

<sup>1</sup> Neglecting interference between the X(3900) and non-resonant continuum.

<sup>2</sup> For  $M^2(\pi^+\pi^-) < 0.65 \text{ GeV}^2$ . Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration.

### X(3900) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 J/\psi\pi$	seen
$\Gamma_2 h_c\pi^\pm$	not seen
$\Gamma_3 \eta_c\pi^+\pi^-$	not seen
$\Gamma_4 (D\overline{D}^*)^\pm$	seen
$\Gamma_5 D^0\overline{D}^{*-} + \text{c.c.}$	seen

See key on page 601

# Meson Particle Listings

## X(3900), X(3915)

$\Gamma_6$	$D^- D^{*0} + \text{c.c.}$	seen
$\Gamma_7$	$\omega \pi^\pm$	not seen
$\Gamma_8$	$J/\psi \eta$	not seen
$\Gamma_9$	$D^+ D^{*-} + \text{c.c.}$	seen
$\Gamma_{10}$	$D^0 \bar{D}^{*0} + \text{c.c.}$	seen

### X(3900) BRANCHING RATIOS

$\Gamma(J/\psi\pi)/\Gamma_{\text{total}}$						$\Gamma_1/\Gamma$	
VALUE	CL%	EVTs	DOCUMENT ID	TECN	CHG	COMMENT	
seen	356		ABLIKIM	15U	BES3	0	$e^+e^- \rightarrow \pi^0\pi^0 J/\psi$
<b>seen</b>	307		ABLIKIM	13T	BES3	$\pm$	$e^+e^- \rightarrow \pi^\pm\pi^\mp J/\psi$
seen	25	<sup>1</sup>	XIAO	13A		0	$4.17 \pi^0 e^+e^- \rightarrow \pi^0\pi^0 J/\psi$

• • • We do not use the following data for averages, fits, limits, etc. • • •

not seen 90 <sup>2</sup>ADOLPH 15D COMP  $\pm$   $\gamma N \rightarrow J/\psi \pi^\pm N$

<sup>1</sup> Obtained by analyzing CLEO-c data but not authored by the CLEO Collaboration.

<sup>2</sup> ADOLPH 15D measure  $B(X(3900)^\pm \rightarrow J/\psi \pi^\pm) \sigma(\gamma N \rightarrow X(3900)^\pm N) / \sigma(\gamma N \rightarrow J/\psi N) < 3.7 \times 10^{-3}$  at 90% CL.

$\Gamma(h_c \pi^\pm)/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
not seen	ABLIKIM	13X	BES3	$\pm$	$e^+e^- \rightarrow h_c \pi^\pm \pi^\mp$

$\Gamma(\eta_c \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
not seen	<sup>1</sup> VINOKUROVA 15	BELL	0	$B^+ \rightarrow K^+ \eta_c \pi^+ \pi^-$	
<sup>1</sup> VINOKUROVA 15 reports $B(B^+ \rightarrow K^+ X(3900)^0) \times B(X \rightarrow \eta_c \pi^+ \pi^-) < 4.7 \times 10^{-5}$ at 90% CL.					

$\Gamma((D\bar{D}^*)^\pm)/\Gamma(J/\psi\pi)$					$\Gamma_4/\Gamma_1$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>6.2 ± 1.1 ± 2.7</b>	<sup>1</sup> ABLIKIM	14A	BES3	$\pm$	$e^+e^- \rightarrow \pi^\pm (D\bar{D}^*)^\mp$
<sup>1</sup> Assuming the same origin of the $(D\bar{D}^*)^\pm$ and $\pi^\pm J/\psi$ decay modes.					

$\Gamma(D^0 D^{*-} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
seen	ABLIKIM	15AC	BES3	$\pm$	$e^+e^- \rightarrow \pi^+ D^0 D^{*-} + \text{c.c.}$
seen	ABLIKIM	14A	BES3	$\pm$	$e^+e^- \rightarrow \pi^+ D^0 D^{*-} + \text{c.c.}$

$\Gamma(D^- D^{*0} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
seen	ABLIKIM	15AC	BES3	$\pm$	$e^+e^- \rightarrow \pi^+ D^- D^{*0} + \text{c.c.}$
seen	ABLIKIM	14A	BES3	$\pm$	$e^+e^- \rightarrow \pi^+ D^- D^{*0} + \text{c.c.}$

$\Gamma(\omega \pi^\pm)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
not seen	ABLIKIM	15R	BES3	$\pm$	$e^+e^- \rightarrow \omega \pi^+ \pi^-$

$\Gamma(J/\psi\eta)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
not seen	ABLIKIM	15Q	BES3	0	$4.0\text{--}4.6 e^+e^- \rightarrow J/\psi \eta \pi^0$

$\Gamma(J/\psi\eta)/\Gamma(J/\psi\pi)$					$\Gamma_8/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.15	90	ABLIKIM	15Q	BES3	0	$4.226\ e^+e^- \rightarrow J/\psi\eta\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.65	90	ABLIKIM	15Q	BES3	0	$4.257\ e^+e^- \rightarrow J/\psi\eta\pi^0$

$\Gamma(D^+ D^{*-} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
seen	ABLIKIM	15AB	BES3	0	$e^+e^- \rightarrow \pi^0 (D\bar{D}^*)^0$

$\Gamma(D^0 \bar{D}^{*0} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{10}/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
seen	ABLIKIM	15AB	BES3	0	$e^+e^- \rightarrow \pi^0 (D\bar{D}^*)^0$

$\Gamma(D^+ D^{*-} + \text{c.c.})/\Gamma(D^0 \bar{D}^{*0} + \text{c.c.})$					$\Gamma_9/\Gamma_{10}$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<b>0.96 ± 0.18 ± 0.12</b>	ABLIKIM	15AB	BES3	0	$e^+e^- \rightarrow \pi^0 (D\bar{D}^*)^0$

### X(3900) REFERENCES

ABLIKIM	15AB	PRL 115 222002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15AC	PR D92 092006	M. Ablikim <i>et al.</i>	(BES III Collab.) JP
ABLIKIM	15Q	PR D92 012008	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15R	PR D92 032009	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	15U	PRL 115 112003	M. Ablikim <i>et al.</i>	(BES III Collab.)
ADOLPH	15D	PL B742 330	C. Adolph <i>et al.</i>	(COMPASS Collab.)
VINOKUROVA	15	JHEP 1506 132	A. Vinokurova <i>et al.</i>	(BELLE Collab.)
ABLIKIM	14A	PRL 112 022001	M. Ablikim <i>et al.</i>	(BES III Collab.) JP
ABLIKIM	13T	PRL 110 252001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM	13X	PRL 111 242001	M. Ablikim <i>et al.</i>	(BES III Collab.)
LIU	13B	PRL 110 252002	Z.Q. Liu <i>et al.</i>	(BELLE Collab.)
XIAO	13A	PL B727 366	T. Xiao <i>et al.</i>	(NWES)

## X(3915)

was  $\chi_{c0}(3915)$ 

$$J^G(J^{PC}) = 0^+(0 \text{ or } 2^{++})$$

The experimental analysis prefers  $J^{PC} = 0^{++}$ . However, a re-analysis presented in ZHOU 15C shows that if helicity-2 dominance assumption is abandoned and a sizable helicity-0 component is allowed, a  $J^{PC} = 2^{++}$  assignment is possible.

### X(3915) MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>3918.4 ± 1.9 OUR AVERAGE</b>				
3919.4 ± 2.2 ± 1.6	59 ± 10	LEES	12AD	BABR $e^+e^- \rightarrow e^+e^- \omega J/\psi$
3919.1 ± 3.8 ± 2.0		DEL-AMO-SA...10B	BABR	$B \rightarrow \omega J/\psi K$
3915 ± 3 ± 2	49 ± 15	UEHARA	10	BELL 10.6 $e^+e^- \rightarrow e^+e^- \omega J/\psi$
3943 ± 11 ± 13	58 ± 11	<sup>1</sup> CHOI	05	BELL $B \rightarrow \omega J/\psi K$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3914.6 ± 3.8 ± 2.0		<sup>1</sup> AUBERT	08W	BABR Superseded by DEL-AMO-SANCHEZ 10B
<sup>1</sup> $\omega J/\psi$ threshold enhancement fitted as an S-wave Breit-Wigner resonance.				

### X(3915) WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>20 ± 5 OUR AVERAGE</b>				Error includes scale factor of 1.1.
13 ± 6 ± 3	59 ± 10	LEES	12AD	BABR $e^+e^- \rightarrow e^+e^- \omega J/\psi$
31 ± 10 ± 8 ± 5		DEL-AMO-SA...10B	BABR	$B \rightarrow \omega J/\psi K$
17 ± 10 ± 3	49 ± 15	UEHARA	10	BELL 10.6 $e^+e^- \rightarrow e^+e^- \omega J/\psi$
87 ± 22 ± 26	58 ± 11	<sup>2</sup> CHOI	05	BELL $B \rightarrow \omega J/\psi K$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
34 ± 12 ± 8 ± 5		<sup>2</sup> AUBERT	08W	BABR Superseded by DEL-AMO-SANCHEZ 10B
<sup>2</sup> $\omega J/\psi$ threshold enhancement fitted as an S-wave Breit-Wigner resonance.				

### X(3915) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\omega J/\psi$	seen
$\Gamma_2$ $\bar{D}^{*0} D^0$	
$\Gamma_3$ $\pi^+ \pi^- \eta_c(1S)$	not seen
$\Gamma_4$ $\eta_c \eta$	not seen
$\Gamma_5$ $\eta_c \pi^0$	not seen
$\Gamma_6$ $K\bar{K}$	not seen
$\Gamma_7$ $\gamma\gamma$	seen

### X(3915) $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(\omega J/\psi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_1\Gamma_7/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>54 ± 9 OUR AVERAGE</b>					
52 ± 10 ± 3	59 ± 10	<sup>3</sup> LEES	12AD	BABR $e^+e^- \rightarrow e^+e^- \omega J/\psi$	
61 ± 17 ± 8	49 ± 15	<sup>3</sup> UEHARA	10	BELL 10.6 $e^+e^- \rightarrow e^+e^- \omega J/\psi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
18 ± 5 ± 2	49 ± 15	<sup>4</sup> UEHARA	10	BELL 10.6 $e^+e^- \rightarrow e^+e^- \omega J/\psi$	
<sup>3</sup> For $J^P = 0^+$ .					
<sup>4</sup> For $J^P = 2^+$ , helicity-2.					

$\Gamma(\pi^+ \pi^- \eta_c(1S)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_3\Gamma_7/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;16</b>	90	LEES	12AE	BABR $e^+e^- \rightarrow e^+e^- \pi^+ \pi^- \eta_c$	

$\Gamma(K\bar{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_6\Gamma_7/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;1.96</b>	90	UEHARA	13	BELL $\gamma\gamma \rightarrow K_S^0 K_S^0$	

### X(3915) BRANCHING RATIOS

$\Gamma(\omega J/\psi)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
seen	<sup>5</sup> DEL-AMO-SA...10B	BABR	$B \rightarrow \omega J/\psi K$		
seen	<sup>6</sup> CHOI	05	BELL $B \rightarrow \omega J/\psi K$		
<sup>5</sup> DEL-AMO-SANCHEZ 10B reports $B(B^\pm \rightarrow X(3915) K^\pm) \times B(X(3915) \rightarrow J/\psi \omega) = (3.0^{+0.7+0.5}_{-0.6-0.3}) \times 10^{-5}$ and $B(B^0 \rightarrow X(3915) K^0) \times B(X(3915) \rightarrow J/\psi \omega) = (2.1 \pm 0.9 \pm 0.3) \times 10^{-5}$ .					
<sup>6</sup> CHOI 05 reports $B(B \rightarrow X(3915) K) \times B(X(3915) \rightarrow J/\psi \omega) = (7.1 \pm 1.3 \pm 3.1) \times 10^{-5}$ .					

# Meson Particle Listings

## X(3915), $\chi_{c2}(2P)$ , X(3940)

### $\Gamma(\omega J/\psi)/\Gamma(\overline{D}^{*0} D^0)$ $\Gamma_1/\Gamma_2$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
>0.71	90	<sup>7</sup> AUSHEV	10	BELL $B \rightarrow \overline{D}^{*0} D^0 K$
<sup>7</sup> By combining the upper limit $B(B \rightarrow X(3915) K) \times B(X(3915) \rightarrow D^{*0} \overline{D}^0) < 0.67 \times 10^{-4}$ from AUSHEV 10 with the average of CHOI 05 and AUBERT 08w measurements $B(B \rightarrow X(3915) K) \times B(X(3915) \rightarrow \omega J/\psi) = (0.51 \pm 0.11) \times 10^{-4}$ .				

### $\Gamma(\eta_c \eta)/\Gamma_{\text{total}}$ $\Gamma_4/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
not seen	<sup>8</sup> VINOKUROVA 15	BELL	$B^+ \rightarrow K^+ \eta_c \eta$
<sup>8</sup> VINOKUROVA 15 reports $B(B^+ \rightarrow K^+ X(3915)^0) \times B(X \rightarrow \eta_c \eta) < 3.3 \times 10^{-5}$ at 90% CL.			

### $\Gamma(\eta_c \pi^0)/\Gamma_{\text{total}}$ $\Gamma_5/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
not seen	<sup>9</sup> VINOKUROVA 15	BELL	$B^+ \rightarrow K^+ \eta_c \pi^0$
<sup>9</sup> VINOKUROVA 15 reports $B(B^+ \rightarrow K^+ X(3915)^0) \times B(X \rightarrow \eta_c \pi^0) < 1.8 \times 10^{-5}$ at 90% CL.			

### $\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ $\Gamma_7/\Gamma$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
seen	59 ± 10	LEES	12AD BABR	$e^+ e^- \rightarrow e^+ e^- \omega J/\psi$
seen		UEHARA	10 BELL	10.6 $e^+ e^- \rightarrow e^+ e^- \omega J/\psi$

#### X(3915) REFERENCES

VINOKUROVA 15	JHEP 1506 132	A. Vinokurova <i>et al.</i>	(BELLE Collab.)
ZHOU 15C	PRL 115 022001	Z.-Y. Zhou, Z. Xiao, H.-Q. Zhou	(BEIJT, NANJ)
UEHARA 13	PTEP 2013 123C01	S. Uehara <i>et al.</i>	(BELLE Collab.)
LEES 12AD	PR D86 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES 12AE	PR D86 092005	J.P. Lees <i>et al.</i>	(BABAR Collab.)
AUSHEV 10	PR D81 031103	T. Aushev <i>et al.</i>	(BELLE Collab.)
DEL-AMO-SA...10B	PR D82 011101	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
UEHARA 10	PRL 104 092001	S. Uehara <i>et al.</i>	(BELLE Collab.)
AUBERT 08W	PRL 101 082001	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHOI 05	PRL 94 182002	S.-K. Choi <i>et al.</i>	(BELLE Collab.)

$\chi_{c2}(2P)$

$I^G(J^{PC}) = 0^+(2^{++})$

#### $\chi_{c2}(2P)$ MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>3927.2 ± 2.6 OUR AVERAGE</b>				
3926.7 ± 2.7 ± 1.1	76 ± 17	AUBERT	10G BABR	10.6 $e^+ e^- \rightarrow e^+ e^- D \overline{D}$
3929 ± 5 ± 2	64	UEHARA	06 BELL	10.6 $e^+ e^- \rightarrow e^+ e^- D \overline{D}$

#### $\chi_{c2}(2P)$ WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>24 ± 6 OUR AVERAGE</b>				
21.3 ± 6.8 ± 3.6	76 ± 17	AUBERT	10G BABR	10.6 $e^+ e^- \rightarrow e^+ e^- D \overline{D}$
29 ± 10 ± 2	64	UEHARA	06 BELL	10.6 $e^+ e^- \rightarrow e^+ e^- D \overline{D}$

#### $\chi_{c2}(2P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\gamma\gamma$	seen
$\Gamma_2$ $K \overline{K} \pi$	
$\Gamma_3$ $K^+ K^- \pi^+ \pi^- \pi^0$	
$\Gamma_4$ $D \overline{D}$	seen
$\Gamma_5$ $D^+ D^-$	seen
$\Gamma_6$ $D^0 \overline{D}^0$	seen
$\Gamma_7$ $\pi^+ \pi^- \eta_c(1S)$	not seen
$\Gamma_8$ $K \overline{K}$	not seen

#### $\chi_{c2}(2P)$ PARTIAL WIDTHS

##### $\chi_{c2}(2P) \Gamma(i) \Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(K\overline{K}\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$				$\Gamma_2\Gamma_1/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<2.1	90	DEL-AMO-SA...11M	BABR	$\gamma\gamma \rightarrow K_S^0 K^\pm \pi^\mp$

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$				$\Gamma_3\Gamma_1/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<3.4	90	DEL-AMO-SA..11M	BABR	$\gamma\gamma \rightarrow K^+ K^- \pi^+ \pi^- \pi^0$

$\Gamma(D\overline{D}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_4\Gamma_1/\Gamma$			
VALUE (keV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.21±0.04 OUR AVERAGE</b>				

<sup>1</sup> Assuming  $B(D^+ D^-) = 0.89 B(D^0 \overline{D}^0)$ .

### $\Gamma(\pi^+ \pi^- \eta_c(1S)) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ $\Gamma_7 \Gamma_1/\Gamma$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<18	90	LEES	12AE BABR	$e^+ e^- \rightarrow e^+ e^- \pi^+ \pi^- \eta_c$

### $\Gamma(K \overline{K}) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$ $\Gamma_8 \Gamma_1/\Gamma$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<0.256	90	UEHARA	13 BELL	$\gamma\gamma \rightarrow K_S^0 K_S^0$

#### $\chi_{c2}(2P)$ BRANCHING RATIOS

### $\Gamma(D^+ D^-)/\Gamma(D^0 \overline{D}^0)$ $\Gamma_5/\Gamma_6$

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.74 ± 0.43 ± 0.16</b>	64	UEHARA	06 BELL	10.6 $e^+ e^- \rightarrow e^+ e^- D \overline{D}$

#### $\chi_{c2}(2P)$ REFERENCES

UEHARA 13	PTEP 2013 123C01	S. Uehara <i>et al.</i>	(BELLE Collab.)
LEES 12AE	PR D86 092005	J.P. Lees <i>et al.</i>	(BABAR Collab.)
DEL-AMO-SA...11M	PR D84 012004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
AUBERT 10G	PR D81 092003	B. Aubert <i>et al.</i>	(BABAR Collab.)
UEHARA 06	PRL 96 082003	S. Uehara <i>et al.</i>	(BELLE Collab.)

X(3940)

$I^G(J^{PC}) = ?^?(?^{??})$

OMITTED FROM SUMMARY TABLE  
Reported by ABE 07, observed in  $e^+ e^- \rightarrow J/\psi X$ .

#### X(3940) MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>3942 ± <sup>7</sup>6 ± 6</b>	52	PAKHLOV	08 BELL	$e^+ e^- \rightarrow J/\psi X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3943 ± 6 ± 6	25	<sup>1</sup> ABE	07 BELL	$e^+ e^- \rightarrow J/\psi X$
3936 ± 14	266	<sup>2</sup> ABE	07 BELL	$e^+ e^- \rightarrow J/\psi(c\overline{c})$
<sup>1</sup> From a fit to $D^{*+} D^-$ and $D^{*0} \overline{D}^0$ events. <sup>2</sup> From the inclusive fit. Not independent of the exclusive measurement by ABE 07.				

#### X(3940) WIDTH

VALUE (MeV)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>37<sup>+26</sup><sub>-15</sub>±8</b>	52	PAKHLOV	08	BELL	$e^+e^- \rightarrow J/\psi X$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<52	90	25	ABE	07	BELL $e^+e^- \rightarrow J/\psi X$

#### X(3940) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D \overline{D}^* + \text{c.c.}$	seen
$\Gamma_2$ $D \overline{D}$	not seen
$\Gamma_3$ $J/\psi \omega$	not seen

#### X(3940) BRANCHING RATIOS

### $\Gamma(D \overline{D}^* + \text{c.c.})/\Gamma_{\text{total}}$ $\Gamma_1/\Gamma$

VALUE	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>0.45	90	25	<sup>3,4</sup> ABE	07 BELL	$e^+ e^- \rightarrow J/\psi X$
<sup>3</sup> For X(3940) decaying to final states with more than two tracks. <sup>4</sup> PAKHLOV 08 finds that the inclusive peak near 3940 MeV/c <sup>2</sup> may consist of several states.					

### $\Gamma(D \overline{D})/\Gamma_{\text{total}}$ $\Gamma_2/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.41	90	<sup>5,6</sup> ABE	07 BELL	$e^+ e^- \rightarrow J/\psi X$
<sup>5</sup> For X(3940) decaying to final states with more than two tracks. <sup>6</sup> PAKHLOV 08 finds that the inclusive peak near 3940 MeV/c <sup>2</sup> may consist of several states.				

### $\Gamma(J/\psi \omega)/\Gamma_{\text{total}}$ $\Gamma_3/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.26	90	<sup>7,8</sup> ABE	07 BELL	$e^+ e^- \rightarrow J/\psi X$
<sup>7</sup> For X(3940) decaying to final states with more than two tracks. <sup>8</sup> PAKHLOV 08 finds that the inclusive peak near 3940 MeV/c <sup>2</sup> may consist of several states.				

#### X(3940) REFERENCES

PAKHLOV 08	PRL 100 202001	P. Pakhlov <i>et al.</i>	(BELLE Collab.)
ABE 07	PRL 98 082001	K. Abe <i>et al.</i>	(BELLE Collab.)

See key on page 601

# Meson Particle Listings

## X(4020), $\psi(4040)$

### X(4020)

$$I(J^P) = 1(?)^?$$

Charged X(4020) seen by ABLIKIM 13X from  $e^+e^- \rightarrow \pi^+\pi^-h_c(1P)$  at c.m. energy from 3.90 to 4.42 GeV as a peak in the invariant mass distribution of the  $\pi^\pm h_c(1P)$  system, and by ABLIKIM 14B from  $e^+e^- \rightarrow (D^*\bar{D}^*)^\pm\pi^\mp$  events in  $(D^*\bar{D}^*)^\pm$  mass. A neutral X(4020) seen by ABLIKIM 14P at three c.m. energies in the same range in  $e^+e^- \rightarrow \pi^0\pi^0h_c(1P)$  as a peak in the larger of the two masses recoiling against a  $\pi^0$ . ABLIKIM 15AA observes a  $5.9\sigma$  signal in  $(D^*\bar{D}^*)^0$  in  $e^+e^- \rightarrow (D^*\bar{D}^*)^0\pi^0$  events using collisions at two c.m. energies. Production rates and mass values support grouping neutral and charged X(4020) together as manifestations of a single  $I = 1$  particle.

### X(4020) MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>4024.1 ± 1.9 OUR AVERAGE</b>					
4025.5 <sup>+2.0</sup> <sub>-4.7</sub> ± 3.1	116	<sup>1</sup> ABLIKIM	15AA	BES3	0 $e^+e^- \rightarrow (D^*\bar{D}^*)^0\pi^0$
4026.3 ± 2.6 ± 3.7	401	<sup>1</sup> ABLIKIM	14B	BES3	± $e^+e^- \rightarrow (D^*\bar{D}^*)^\pm\pi^\mp$
4023.9 ± 2.2 ± 3.8	61	<sup>1,2</sup> ABLIKIM	14P	BES3	0 $e^+e^- \rightarrow \pi^0\pi^0h_c$
4022.9 ± 0.8 ± 2.7	253	<sup>1</sup> ABLIKIM	13X	BES3	± $e^+e^- \rightarrow \pi^+\pi^-\pi^0h_c$

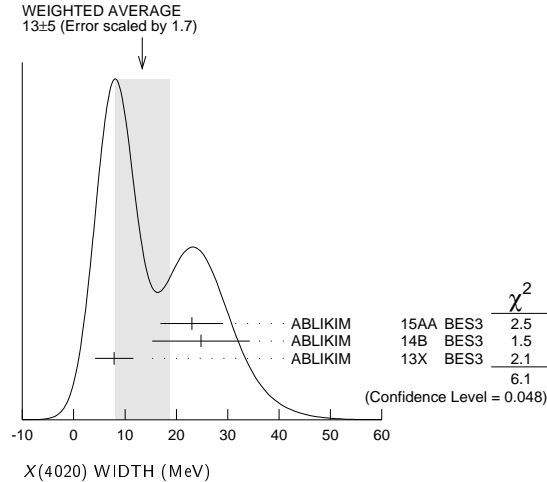
<sup>1</sup> Neglecting interference between the X(4020) and non-resonant continuum.

<sup>2</sup> Assuming  $J^P = 1^+$  and width of  $7.9 \pm 2.6$  MeV.

### X(4020) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>13 ± 5 OUR AVERAGE</b>					Error includes scale factor of 1.7. See the ideogram below.
23.0 ± 6.0 ± 1.0	116	<sup>1</sup> ABLIKIM	15AA	BES3	0 $e^+e^- \rightarrow (D^*\bar{D}^*)^0\pi^0$
24.8 ± 5.6 ± 7.7	401	<sup>1</sup> ABLIKIM	14B	BES3	± $e^+e^- \rightarrow (D^*\bar{D}^*)^\pm\pi^\mp$
7.9 ± 2.7 ± 2.6	253	<sup>1</sup> ABLIKIM	13X	BES3	± $e^+e^- \rightarrow \pi^+\pi^-\pi^0h_c$

<sup>1</sup> Neglecting interference between the X(4020) and non-resonant continuum.



### X(4020) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $h_c(1P)\pi$	seen
$\Gamma_2$ $D^*\bar{D}^*$	seen
$\Gamma_3$ $D\bar{D}^* + \text{c.c.}$	not seen
$\Gamma_4$ $\eta_c\pi^+\pi^-$	not seen

### X(4020) BRANCHING RATIOS

$\Gamma(h_c(1P)\pi)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_1/\Gamma$
seen	61		ABLIKIM	14P	BES3	0 $e^+e^- \rightarrow \pi^0\pi^0h_c$	
seen	253		ABLIKIM	13X	BES3	± $e^+e^- \rightarrow \pi^+\pi^-\pi^0h_c$	

$\Gamma(D^*\bar{D}^*)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_2/\Gamma$
seen	116		<sup>1</sup> ABLIKIM	15AA	BES3	0 $e^+e^- \rightarrow (D^*\bar{D}^*)^0\pi^0$	
seen	401		<sup>1</sup> ABLIKIM	14B	BES3	± $e^+e^- \rightarrow (D^*\bar{D}^*)^\pm\pi^\mp$	

<sup>1</sup> Neglecting interference between the X(4020) and non-resonant continuum.

$\Gamma(D\bar{D}^* + \text{c.c.})/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_3/\Gamma$
not seen		ABLIKIM	15AC	BES3	± $e^+e^- \rightarrow \pi^\pm(D\bar{D}^*)^\mp$	

$\Gamma(\eta_c\pi^+\pi^-)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
not seen		<sup>1</sup> VINOKUROVA 15	BELL	$B^+ \rightarrow K^+\eta_c\pi^+\pi^-$	
		<sup>1</sup> VINOKUROVA 15 reports $B(B^+ \rightarrow K^+X(4020)^0) \times B(X \rightarrow \eta_c\pi^+\pi^-) < 1.6 \times 10^{-5}$ at 90% CL.			

### X(4020) REFERENCES

ABLIKIM 15AA	PRL 115 182002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM 15AC	PR D92 092006	M. Ablikim <i>et al.</i>	(BES III Collab.)
VINOKUROVA 15	JHEP 1506 132	A. Vinokurova <i>et al.</i>	(Belle Collab.)
ABLIKIM 14B	PRL 112 132001	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM 14P	PRL 113 212002	M. Ablikim <i>et al.</i>	(BES III Collab.)
ABLIKIM 13X	PRL 111 242001	M. Ablikim <i>et al.</i>	(BES III Collab.)

### $\psi(4040)$

$$I^G(J^{PC}) = 0^-(1^{--})$$

### $\psi(4040)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4039 ± 1 OUR ESTIMATE</b>			
<b>4039.6 ± 4.3</b>	<sup>1</sup> ABLIKIM	08D	BES2 $e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4034 ± 6	<sup>2</sup> MO	10	RVUE $e^+e^- \rightarrow \text{hadrons}$
4037 ± 2	<sup>3</sup> SETH	05A	RVUE $e^+e^- \rightarrow \text{hadrons}$
4040 ± 1	<sup>4</sup> SETH	05A	RVUE $e^+e^- \rightarrow \text{hadrons}$
4040 ± 10	BRANDELIK	78C	DASP $e^+e^-$

<sup>1</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (130 \pm 46)^\circ$ .

<sup>2</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects.

<sup>3</sup> From a fit to Crystal Ball (OSTERHELD 86) data.

<sup>4</sup> From a fit to BES (BAI 02c) data.

### $\psi(4040)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>80 ± 10 OUR ESTIMATE</b>			
<b>84.5 ± 12.3</b>	<sup>5</sup> ABLIKIM	08D	BES2 $e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
87 ± 11	<sup>6</sup> MO	10	RVUE $e^+e^- \rightarrow \text{hadrons}$
85 ± 10	<sup>7</sup> SETH	05A	RVUE $e^+e^- \rightarrow \text{hadrons}$
89 ± 6	<sup>8</sup> SETH	05A	RVUE $e^+e^- \rightarrow \text{hadrons}$
52 ± 10	BRANDELIK	78C	DASP $e^+e^-$

<sup>5</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (130 \pm 46)^\circ$ .

<sup>6</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects.

<sup>7</sup> From a fit to Crystal Ball (OSTERHELD 86) data.

<sup>8</sup> From a fit to BES (BAI 02c) data.

### $\psi(4040)$ DECAY MODES

Due to the complexity of the  $c\bar{c}$  threshold region, in this listing, “seen” (“not seen”) means that a cross section for the mode in question has been measured at effective  $\sqrt{s}$  near this particle's central mass value, more (less) than  $2\sigma$  above zero, without regard to any peaking behavior in  $\sqrt{s}$  or absence thereof. See mode listing(s) for details and references.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $e^+e^-$	$(1.07 \pm 0.16) \times 10^{-5}$	
$\Gamma_2$ $D\bar{D}$	seen	
$\Gamma_3$ $D^0\bar{D}^0$	seen	
$\Gamma_4$ $D^+D^-$	seen	
$\Gamma_5$ $D^*\bar{D} + \text{c.c.}$	seen	
$\Gamma_6$ $D^*(2007)^0\bar{D}^0 + \text{c.c.}$	seen	
$\Gamma_7$ $D^*(2010)^+D^- + \text{c.c.}$	seen	
$\Gamma_8$ $D^*\bar{D}^*$	seen	
$\Gamma_9$ $D^*(2007)^0\bar{D}^*(2007)^0$	seen	
$\Gamma_{10}$ $D^*(2010)^+D^*(2010)^-$	seen	
$\Gamma_{11}$ $D\bar{D}\pi$ (excl. $D^*\bar{D}$ )		
$\Gamma_{12}$ $D^0D^-\pi^+ + \text{c.c.}$ (excl. $D^*(2007)^0\bar{D}^0 + \text{c.c.}$ , $D^*(2010)^+D^- + \text{c.c.}$ )	not seen	

# Meson Particle Listings

## $\psi(4040)$

$\Gamma_{13}$	$D\overline{D}^*\pi$ (excl. $D^*\overline{D}^*$ )	not seen
$\Gamma_{14}$	$D^0\overline{D}^{*-}\pi^++\text{c.c.}$ (excl. $D^*(2010)^+\overline{D}^*(2010)^-$ )	seen
$\Gamma_{15}$	$D_s^+D_s^-$	seen
$\Gamma_{16}$	$J/\psi(1S)$ hadrons	
$\Gamma_{17}$	$J/\psi\pi^+\pi^-$	$< 4 \times 10^{-3}$
$\Gamma_{18}$	$J/\psi\pi^0\pi^0$	$< 2 \times 10^{-3}$
$\Gamma_{19}$	$J/\psi\eta$	$(5.2 \pm 0.7) \times 10^{-3}$
$\Gamma_{20}$	$J/\psi\pi^0$	$< 2.8 \times 10^{-4}$
$\Gamma_{21}$	$J/\psi\pi^+\pi^-\pi^0$	$< 2 \times 10^{-3}$
$\Gamma_{22}$	$\chi_{c1}\gamma$	$< 3.4 \times 10^{-3}$
$\Gamma_{23}$	$\chi_{c2}\gamma$	$< 5 \times 10^{-3}$
$\Gamma_{24}$	$\chi_{c1}\pi^+\pi^-\pi^0$	$< 1.1 \%$
$\Gamma_{25}$	$\chi_{c2}\pi^+\pi^-\pi^0$	$< 3.2 \%$
$\Gamma_{26}$	$h_c(1P)\pi^+\pi^-$	$< 3 \times 10^{-3}$
$\Gamma_{27}$	$\phi\pi^+\pi^-$	$< 3 \times 10^{-3}$
$\Gamma_{28}$	$\Lambda\overline{\Lambda}\pi^+\pi^-$	$< 2.9 \times 10^{-4}$
$\Gamma_{29}$	$\Lambda\overline{\Lambda}\pi^0$	$< 9 \times 10^{-5}$
$\Gamma_{30}$	$\Lambda\overline{\Lambda}\eta$	$< 3.0 \times 10^{-4}$
$\Gamma_{31}$	$\Sigma^+\Sigma^-$	$< 1.3 \times 10^{-4}$
$\Gamma_{32}$	$\Sigma^0\Sigma^0$	$< 7 \times 10^{-5}$
$\Gamma_{33}$	$\Xi^+\Xi^-$	$< 1.6 \times 10^{-4}$
$\Gamma_{34}$	$\Xi^0\Xi^0$	$< 1.8 \times 10^{-4}$
$\Gamma_{35}$	$\mu^+\mu^-$	

### $\psi(4040)$ PARTIAL WIDTHS

$\Gamma(e^+e^-)$				$\Gamma_1$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
<b><math>0.86 \pm 0.07</math> OUR ESTIMATE</b>				
<b><math>0.83 \pm 0.20</math></b>	<sup>9</sup> ABLIKIM	08D	BES2	$e^+e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.6 \pm 1.4$	<sup>10</sup> MO	10	RVUE	$e^+e^- \rightarrow$ hadrons
$0.88 \pm 0.11$	<sup>11</sup> SETH	05A	RVUE	$e^+e^- \rightarrow$ hadrons
$0.91 \pm 0.13$	<sup>12</sup> SETH	05A	RVUE	$e^+e^- \rightarrow$ hadrons
$0.75 \pm 0.15$	BRA NDELIK	78C	DASP	$e^+e^-$
<sup>9</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the $\psi(3770)$ , $\psi(4040)$ , $\psi(4160)$ , and $\psi(4415)$ resonances. Phase angle fixed in the fit to $\delta = (130 \pm 46)^\circ$ .				
<sup>10</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the $\psi(4040)$ , $\psi(4160)$ and $\psi(4415)$ resonances and including interference effects. Four sets of solutions are obtained with the same fit quality, mass and total width, but with different $e^+e^-$ partial widths. We quote only the range of values.				
<sup>11</sup> From a fit to Crystal Ball (OSTERHELD 86) data.				
<sup>12</sup> From a fit to BES (BAI 02c) data.				

### $\psi(4040) \Gamma(i) \times \Gamma(e^+e^-)/\Gamma(\text{total})$

$\Gamma(\chi_{c1}\gamma) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_{22}\Gamma_1/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2.9</math></b>	90	<sup>13</sup> HAN	15	BELL
<sup>13</sup> Using $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.21)\%$ .				

$\Gamma(\chi_{c2}\gamma) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_{23}\Gamma_1/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4.6</math></b>	90	<sup>14</sup> HAN	15	BELL
<sup>14</sup> Using $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.21)\%$ .				

### $\psi(4040) \Gamma(i) \times \Gamma(e^+e^-)/\Gamma^2(\text{total})$

$\Gamma(J/\psi\eta)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_{19}/\Gamma \times \Gamma_1/\Gamma$
VALUE (units $10^{-8}$ )	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$5.1 \pm 1.4 \pm 1.5$	<sup>15</sup> WANG	13B	BELL	$e^+e^- \rightarrow J/\psi\eta\gamma$
$12.8 \pm 2.1 \pm 1.9$	<sup>16</sup> WANG	13B	BELL	$e^+e^- \rightarrow J/\psi\eta\gamma$
<sup>15</sup> Solution I of two equivalent solutions in a fit using two interfering resonances. Mass and width fixed at 4039 MeV and 80 MeV, respectively.				
<sup>16</sup> Solution II of two equivalent solutions in a fit using two interfering resonances. Mass and width fixed at 4039 MeV and 80 MeV, respectively.				

### $\psi(4040)$ BRANCHING RATIOS

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (units $10^{-9}$ )	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$\sim 1.0$	FELDMAN	77	MRK1	$e^+e^-$

$\Gamma(D^0\overline{D}^0)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	09M	BABR	$e^+e^- \rightarrow D^0\overline{D}^0\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D^0\overline{D}^0$
seen	PAKHLOVA	08	BELL	$e^+e^- \rightarrow D^0\overline{D}^0\gamma$

$\Gamma(D^+D^-)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	09M	BABR	$e^+e^- \rightarrow D^+D^-\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D^+D^-$
seen	PAKHLOVA	08	BELL	$e^+e^- \rightarrow D^+D^-\gamma$

$\Gamma(D\overline{D})/\Gamma(D^*\overline{D}^+ + \text{c.c.})$				$\Gamma_2/\Gamma_5$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.24 \pm 0.05 \pm 0.12</math></b>	AUBERT	09M	BABR	$e^+e^- \rightarrow \gamma D^{(*)}\overline{D}$

$\Gamma(D^0\overline{D}^0)/\Gamma(D^*(2007)^0\overline{D}^0 + \text{c.c.})$				$\Gamma_3/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.05 \pm 0.03</math></b>	<sup>17</sup> GOLDHABER	77	MRK1	$e^+e^-$
<sup>17</sup> Phase-space factor ( $p^3$ ) explicitly removed.				

$\Gamma(D^*(2007)^0\overline{D}^0 + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	09M	BABR	$e^+e^- \rightarrow D^{*0}\overline{D}^0\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D^{*0}\overline{D}^0$

$\Gamma(D^*(2010)^+\overline{D}^- + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	09M	BABR	$e^+e^- \rightarrow D^{*+}D^-\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D^{*+}D^-$
seen	PAKHLOVA	07	BELL	$e^+e^- \rightarrow D^{*+}D^-\gamma$

$\Gamma(D^*(2010)^+\overline{D}^- + \text{c.c.})/\Gamma(D^*(2007)^0\overline{D}^0 + \text{c.c.})$				$\Gamma_7/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.95 \pm 0.09 \pm 0.10</math></b>	AUBERT	09M	BABR	$e^+e^- \rightarrow \gamma D^*\overline{D}$

$\Gamma(D^*\overline{D}^*)/\Gamma(D^*\overline{D}^+ + \text{c.c.})$				$\Gamma_8/\Gamma_5$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.18 \pm 0.14 \pm 0.03</math></b>	AUBERT	09M	BABR	$e^+e^- \rightarrow \gamma D^{(*)}\overline{D}^{(*)}$

$\Gamma(D^*(2007)^0\overline{D}^*(2007)^0)/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	09M	BABR	$e^+e^- \rightarrow D^{*0}\overline{D}^{*0}\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D^{*0}\overline{D}^{*0}$

$\Gamma(D^*(2007)^0\overline{D}^*(2007)^0)/\Gamma(D^*(2007)^0\overline{D}^0 + \text{c.c.})$				$\Gamma_9/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>32.0 \pm 12.0</math></b>	<sup>18</sup> GOLDHABER	77	MRK1	$e^+e^-$
<sup>18</sup> Phase-space factor ( $p^3$ ) explicitly removed.				

$\Gamma(D^*(2010)^+\overline{D}^*(2010)^-)/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AUBERT	09M	BABR	$e^+e^- \rightarrow D^{*+}D^{*-}\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D^{*+}D^{*-}$
seen	PAKHLOVA	07	BELL	$e^+e^- \rightarrow D^{*+}D^{*-}\gamma$

$\Gamma(D^0\overline{D}^-\pi^+ + \text{c.c. (excl. } D^*(2007)^0\overline{D}^0 + \text{c.c., } D^*(2010)^+\overline{D}^- + \text{c.c.}))/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
not seen	PAKHLOVA	08A	BELL	$e^+e^- \rightarrow D^0\overline{D}^-\pi^+\gamma$

$\Gamma(D\overline{D}^*\pi \text{ (excl. } D^*\overline{D}^*))/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
not seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D\overline{D}^*\pi$

$\Gamma(D^0\overline{D}^{*-}\pi^+ + \text{c.c. (excl. } D^*(2010)^+\overline{D}^*(2010)^-))/\Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	PAKHLOVA	09	BELL	$e^+e^- \rightarrow D^0\overline{D}^{*-}\pi^+\gamma$

$\Gamma(D_s^+D_s^-)/\Gamma_{\text{total}}$				$\Gamma_{15}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	PAKHLOVA	11	BELL	$e^+e^- \rightarrow D_s^+D_s^-\gamma$
seen	DEL-AMO-SA...10N	BABR		$e^+e^- \rightarrow D_s^+D_s^-\gamma$
seen	CRONIN-HEN..09	CLEO		$e^+e^- \rightarrow D_s^+D_s^-$

$\Gamma(J/\psi\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{17}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 4</math></b>	90	COAN	06	CLEO
$3.97\text{--}4.06\text{ }e^+e^- \rightarrow$ hadrons				

$\Gamma(J/\psi\pi^0\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{18}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 2</math></b>	90	COAN	06	CLEO
$3.97\text{--}4.06\text{ }e^+e^- \rightarrow$ hadrons				

See key on page 601

# Meson Particle Listings

## $\psi(4040), X(4050)^\pm$

### $\Gamma(J/\psi\eta)/\Gamma_{\text{total}}$ $\Gamma_{19}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>5.2 \pm 0.5 \pm 0.5</math></b>		<sup>19</sup> ABLIKIM	12K BES3	$e^+e^- \rightarrow \ell^+\ell^-2\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<7	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$
<sup>19</sup> ABLIKIM 12K measure $\sigma(e^+e^- \rightarrow J/\psi\eta) = 32.1 \pm 2.8 \pm 1.3\text{ pb}$ . They assume the $\eta J/\psi$ fully originates from $\psi(4040)$ decays.				

### $\Gamma(J/\psi\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{20}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.28</b>	90	<sup>20</sup> ABLIKIM	12K BES3	$e^+e^- \rightarrow \ell^+\ell^-2\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$
<sup>20</sup> ABLIKIM 12K measure $\sigma(e^+e^- \rightarrow J/\psi\pi^0) < 1.6\text{ pb}$ . They assume the $\eta J/\psi$ fully originates from $\psi(4040)$ decays.				

### $\Gamma(J/\psi\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{21}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2</b>	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(\chi_{c1}\gamma)/\Gamma_{\text{total}}$ $\Gamma_{22}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<11	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(\chi_{c2}\gamma)/\Gamma_{\text{total}}$ $\Gamma_{23}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<17	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(\chi_{c1}\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{24}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;11</b>	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(\chi_{c2}\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{25}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;32</b>	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(h_c(1P)\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{26}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3</b>	90	<sup>21</sup> PEDLAR	11 CLEO	$e^+e^- \rightarrow h_c(1P)\pi^+\pi^-$
<sup>21</sup> From several values of $\sqrt{s}$ near the peak of the $\psi(4040)$ , PEDLAR 11 measures $\sigma(e^+e^- \rightarrow h_c(1P)\pi^+\pi^-) = 1.0 \pm 8.0 \pm 5.4 \pm 0.2\text{ pb}$ , where the errors are statistical, systematic, and due to uncertainty in $B(\psi(2S) \rightarrow \pi^0 h_c(1P))$ , respectively.				

### $\Gamma(\phi\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{27}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3</b>	90	COAN	06 CLEO	$3.97\text{--}4.06\ e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(\Lambda\bar{\Lambda}\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{28}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.9</b>	90	<sup>22</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>22</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\Gamma(\Lambda\bar{\Lambda}\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{29}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.9</b>	90	<sup>23</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>23</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\Gamma(\Lambda\bar{\Lambda}\eta)/\Gamma_{\text{total}}$ $\Gamma_{30}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.0</b>	90	<sup>24</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>24</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\Gamma(\Sigma^+\bar{\Sigma}^-)/\Gamma_{\text{total}}$ $\Gamma_{31}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.3</b>	90	<sup>25</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>25</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\Gamma(\Sigma^0\bar{\Sigma}^0)/\Gamma_{\text{total}}$ $\Gamma_{32}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.7</b>	90	<sup>26</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>26</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\Gamma(\Xi^+\bar{\Xi}^-)/\Gamma_{\text{total}}$ $\Gamma_{33}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.6</b>	90	<sup>27</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>27</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\Gamma(\Xi^0\bar{\Xi}^0)/\Gamma_{\text{total}}$ $\Gamma_{34}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.8</b>	90	<sup>28</sup> ABLIKIM	13Q BES3	$e^+e^- \rightarrow \psi(4040)$
<sup>28</sup> Assuming that interference effects between resonance and continuum can be neglected.				

### $\psi(4040)$ REFERENCES

HAN	15	PR D92 012011	Y.L. Han <i>et al.</i>	(BELLE Collab.)
ABLIKIM	13Q	PR D87 112011	Ablikim M. <i>et al.</i>	(BES III Collab.)
WANG	13B	PR D87 051101	X.L. Wang <i>et al.</i>	(BELLE Collab.)
ABLIKIM	12K	PR D86 071101	M. Ablikim <i>et al.</i>	(BES III Collab.)
PAKHLOVA	11	PR D83 011101	G. Pakhlova <i>et al.</i>	(CLEO Collab.)
PEDLAR	11	PRL 107 041803	T. Pedlar <i>et al.</i>	(CLEO Collab.)
DEL-AMO-SA...	10N	PR D82 052004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
MO	10	PR D82 077501	X.H. Mo, C.Z. Yuan, P. Wang	(BHEP)
AUBERT	09M	PR D79 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)
CRONIN-HEN...	09	PR D80 072001	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
PAKHLOVA	09	PR D80 091101	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
ABLIKIM	08D	PL B660 315	M. Ablikim <i>et al.</i>	(BES Collab.)
PAKHLOVA	08	PR D77 011103	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PAKHLOVA	08A	PRL 100 062001	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PAKHLOVA	07	PRL 98 092001	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
COAN	06	PRL 96 162003	T.E. Coan <i>et al.</i>	(CLEO Collab.)
SETH	05A	PR D72 017501	K.K. Seth	
BAI	02C	PRL 88 101802	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	00	PRL 84 594	J.Z. Bai <i>et al.</i>	(BES Collab.)
OSTERHELD	86	SLAC-PUB-4160	A. Osterheld <i>et al.</i>	(SLAC Crystal Ball Collab.)
BRANDELIK	78C	PL 76B 361	R. Brandelik <i>et al.</i>	(DASP Collab.)
Also		ZPHY C1 233	R. Brandelik <i>et al.</i>	(DASP Collab.)
FELDMAN	77	PRPL 33C 285	G.J. Feldman, M.L. Perl	(LBL, SLAC)
GOLDBABER	77	PL 69B 503	G. Goldhaber <i>et al.</i>	(Mark I Collab.)

## $X(4050)^\pm$

$$I(J^P) = ?(??)$$

### OMITTED FROM SUMMARY TABLE

Observed by MIZUK 08 in the  $\pi^+\chi_{c1}(1P)$  invariant mass distribution in  $\bar{B}^0 \rightarrow K^-\pi^+\chi_{c1}(1P)$  decays. Not seen by LEES 12B in this same mode after accounting for  $K\pi$  resonant mass and angular structure.

### $X(4050)^\pm$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>4051 \pm 14^{+20}_{-41}</math></b>	<sup>1</sup> MIZUK	08 BELL	$\bar{B}^0 \rightarrow K^-\pi^+\chi_{c1}(1P)$

<sup>1</sup> From a Dalitz plot analysis with two Breit-Wigner amplitudes.

### $X(4050)^\pm$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>82^{+21+47}_{-17-22}</math></b>	<sup>1</sup> MIZUK	08 BELL	$\bar{B}^0 \rightarrow K^-\pi^+\chi_{c1}(1P)$

<sup>1</sup> From a Dalitz plot analysis with two Breit-Wigner amplitudes.

### $X(4050)^\pm$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1\ \pi^+\chi_{c1}(1P)$	seen

### $X(4050)^\pm$ BRANCHING RATIOS

### $\Gamma(\pi^+\chi_{c1}(1P))/\Gamma_{\text{total}}$ $\Gamma_1/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
seen	<sup>1</sup> MIZUK	08 BELL	$\bar{B}^0 \rightarrow K^-\pi^+\chi_{c1}(1P)$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	<sup>2</sup> LEES	12B BABR	$B \rightarrow K\pi\chi_{c1}(1P)$

<sup>1</sup> With a product branching fraction measurement of  $B(\bar{B}^0 \rightarrow K^-\chi_{c1}(4050)^+) \times B(X(4050)^+ \rightarrow \pi^+\chi_{c1}(1P)) = (3.0^{+1.5+3.7}_{-0.8-1.6}) \times 10^{-5}$ .

<sup>2</sup> With a product branching fraction limit of  $B(\bar{B}^0 \rightarrow X(4050)^+K^-) \times B(X(4050)^+ \rightarrow \chi_{c1}\pi^+) < 1.8 \times 10^{-5}$  at 90% CL.

### $X(4050)^\pm$ REFERENCES

LEES	12B	PR D85 052003	J.P. Lees <i>et al.</i>	(BABAR Collab.)
MIZUK	08	PR D78 072004	R. Mizuk <i>et al.</i>	(BELLE Collab.)

Meson Particle Listings

X(4055)<sup>±</sup>, X(4140)

X(4055)<sup>±</sup>

$I(J^P) = ?(?^?)$

OMITTED FROM SUMMARY TABLE  
Needs confirmation. Seen by WANG 15A in the  $\psi(2S)\pi^+\pi^-$  invariant mass distribution in  $X(4360) \rightarrow \psi(2S)\pi^+\pi^-$  decay.

X(4055) <sup>±</sup> MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
4054±3±1	<sup>1</sup> WANG	15A	BELL	10.58 e <sup>+</sup> e <sup>-</sup> → $\gamma\pi^+\pi^-\psi(2S)$
<sup>1</sup> Statistical significance of 3.5 σ.				

X(4055) <sup>±</sup> WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
45±11±6	<sup>1</sup> WANG	15A	BELL	10.58 e <sup>+</sup> e <sup>-</sup> → $\gamma\pi^+\pi^-\psi(2S)$
<sup>1</sup> Statistical significance of 3.5 σ.				

X(4055) <sup>±</sup> DECAY MODES		
Mode	Fraction (Γ <sub>i</sub> /Γ)	
Γ <sub>1</sub> π <sup>+</sup> ψ(2S)	seen	

X(4055) <sup>±</sup> BRANCHING RATIOS			
Γ(π <sup>+</sup> ψ(2S))/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT
seen	<sup>1</sup> WANG	15A	BELL
<sup>1</sup> Statistical significance of 3.5 σ.			

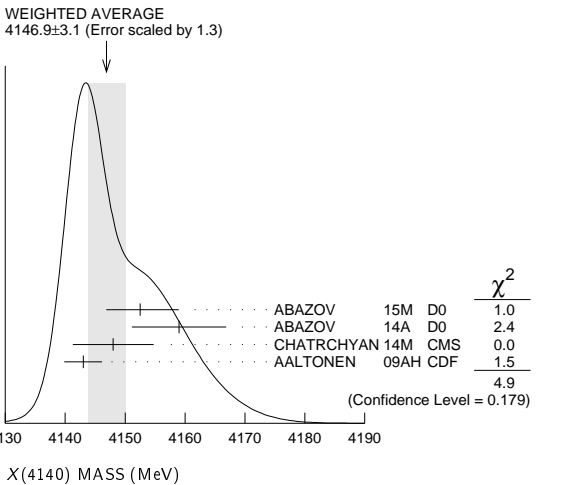
X(4055) <sup>±</sup> REFERENCES			
WANG	15A	PR D91 112007	X.L. Wang <i>et al.</i> (BELLE Collab.)

X(4140)

$I^G(J^{PC}) = 0^+(?^{?+})$

Seen by AALTONEN 09AH, ABAZOV 14A, CHATRCHYAN 14M in  $B^+ \rightarrow X K^+$ ,  $X \rightarrow J/\psi\phi$ , and by ABAZOV 15M separately in both prompt (4.7 σ) and non-prompt (5.6 σ) production in  $p\bar{p} \rightarrow J/\psi\phi + \text{anything}$ . Not seen by SHEN 10 in  $\gamma\gamma \rightarrow J/\psi\phi$ , AAIJ 12AA in  $B^+ \rightarrow J/\psi\phi K^+$ , and ABLIKIM 15 in  $e^+e^- \rightarrow \gamma J/\psi\phi$  at  $\sqrt{s} = 4.23, 4.26, 4.36$  GeV.

X(4140) MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
4146.9±3.1 OUR AVERAGE		Error includes scale factor of 1.3. See the ideogram below.		
4152.5±1.7+6.2-5.4	616	<sup>1</sup> ABAZOV	15M D0	$p\bar{p} \rightarrow J/\psi\phi + \text{anything}$
4159.0±4.3±6.6	52	<sup>2</sup> ABAZOV	14A D0	$B^+ \rightarrow J/\psi\phi K^+$
4148.0±2.4±6.3	0.3k	<sup>3</sup> CHATRCHYAN	14M CMS	$B^+ \rightarrow J/\psi\phi K^+$
4143.0±2.9±1.2	14	<sup>4</sup> AALTONEN	09AH CDF	$B^+ \rightarrow J/\psi\phi K^+$



- <sup>1</sup> Statistical significance of more than 6 σ.
- <sup>2</sup> Statistical significance of 3.1 σ.
- <sup>3</sup> From a fit assuming an S-wave relativistic Breit-Wigner shape above a three-body phase-space non-resonant component with statistical significance of more than 5 σ.
- <sup>4</sup> Statistical significance of 3.8 σ.

X(4140) WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
15 ± 6 OUR AVERAGE				
16.3± 5.6±11.4	616	<sup>1</sup> ABAZOV	15M D0	$p\bar{p} \rightarrow J/\psi\phi + \text{anything}$
20 ±13 ± 3/8	52	<sup>2</sup> ABAZOV	14A D0	$B^+ \rightarrow J/\psi\phi K^+$
28 +15-11 ±19	0.3k	<sup>3</sup> CHATRCHYAN	14M CMS	$B^+ \rightarrow J/\psi\phi K^+$
11.7± 8.3/5.0 ± 3.7	14	<sup>4</sup> AALTONEN	09AH CDF	$B^+ \rightarrow J/\psi\phi K^+$
<sup>1</sup> Statistical significance of more than 6 σ.				
<sup>2</sup> Statistical significance of 3.1 σ.				
<sup>3</sup> From a fit assuming an S-wave relativistic Breit-Wigner shape above a three-body phase-space non-resonant component with statistical significance of more than 5 σ.				
<sup>4</sup> Statistical significance of 3.8 σ.				

X(4140) DECAY MODES		
Mode	Fraction (Γ <sub>i</sub> /Γ)	
Γ <sub>1</sub> J/ψφ	seen	
Γ <sub>2</sub> γγ	not seen	

X(4140) Γ(i)Γ(γγ)/Γ(total)			
Γ(γγ) × Γ(J/ψφ)/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT
VALUE (eV) CL% <41 90	<sup>1</sup> SHEN	10	BELL
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 6 90	<sup>2</sup> SHEN	10	BELL
<sup>1</sup> For J <sup>P</sup> = 0 <sup>+</sup> . <sup>2</sup> For J <sup>P</sup> = 2 <sup>+</sup> .			

X(4140) BRANCHING RATIOS				
Γ(J/ψφ)/Γ <sub>total</sub>	EVTS	DOCUMENT ID	TECN	COMMENT
seen	616	<sup>1</sup> ABAZOV	15M D0	$p\bar{p} \rightarrow J/\psi\phi + \text{anything}$
seen	52	<sup>2</sup> ABAZOV	14A D0	$B^+ \rightarrow J/\psi\phi K^+$
seen	0.3k	<sup>3</sup> CHATRCHYAN	14M CMS	$B^+ \rightarrow J/\psi\phi K^+$
seen	14	<sup>4</sup> AALTONEN	09AH CDF	$B^+ \rightarrow J/\psi\phi K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen		<sup>5</sup> ABLIKIM	15 BES3	$e^+e^- \rightarrow \gamma\phi J/\psi$
not seen		<sup>6</sup> AAIJ	12AA LHCb	$p\bar{p} \rightarrow B^+ X$ at 7 TeV

- <sup>1</sup> Statistical significance of more than 6 σ.
- <sup>2</sup> ABAZOV 14A reports  $B(B^+ \rightarrow X(4140) K^+)/B(B^+ \rightarrow J/\psi\phi K^+) = (19 \pm 7 \pm 4)\%$  with 3.1 σ significance.
- <sup>3</sup> From a fit assuming an S-wave relativistic Breit-Wigner shape above a three-body phase-space non-resonant component with statistical significance of more than 5 σ.
- <sup>4</sup> Statistical significance of 3.8 σ.
- <sup>5</sup> Reported  $\sigma(e^+e^- \rightarrow \gamma X(4140)) \cdot B(X(4140) \rightarrow J/\psi\phi) < 0.35, 0.28$ , and  $0.33$  pb at 4.23, 4.26, and 4.36 GeV, respectively, at 90% CL.
- <sup>6</sup> Reported  $B(B^+ \rightarrow X(4140) K^+) \cdot B(X(4140) \rightarrow J/\psi\phi)/B(B^+ \rightarrow J/\psi\phi K^+) < 0.07$  at 90% CL.

Γ(γγ)/Γ <sub>total</sub>			
VALUE	DOCUMENT ID	TECN	COMMENT
not seen	SHEN	10	BELL
10.6 e <sup>+</sup> e <sup>-</sup> → e <sup>+</sup> e <sup>-</sup> J/ψφ			

X(4140) REFERENCES			
ABAZOV	15M	PRL 115 232001	V.M. Abazov <i>et al.</i> (D0 Collab.)
ABLIKIM	15	PR D91 032002	M. Ablikim <i>et al.</i> (BES III Collab.)
ABAZOV	14A	PR D89 012004	V.M. Abazov <i>et al.</i> (D0 Collab.)
CHATRCHYAN	14M	PL B734 261	S. Chatrchyan <i>et al.</i> (CMS Collab.)
AAIJ	12AA	PR D85 091103	R. Aaij <i>et al.</i> (LHCb Collab.)
SHEN	10	PRL 104 112004	C.P. Shen <i>et al.</i> (BELLE Collab.)
AALTONEN	09AH	PRL 102 242002	T. Aaltonen <i>et al.</i> (CDF Collab.)



$\psi(4160)$

$$\psi(4160)$$

$$J^G(J^{PC}) = 0^-(1^{--})$$

$\psi(4160)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4191 ± 5 OUR AVERAGE</b>			
4191 ± 9	AAIJ	13bc LHCB	$B^+ \rightarrow K^+ \mu^+ \mu^-$
4191.7 ± 6.5	<sup>1</sup> ABLIKIM	08d BES2	$e^+ e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4193 ± 7	<sup>2</sup> MO	10 RVUE	$e^+ e^- \rightarrow$ hadrons
4151 ± 4	<sup>3</sup> SETH	05A RVUE	$e^+ e^- \rightarrow$ hadrons
4155 ± 5	<sup>4</sup> SETH	05A RVUE	$e^+ e^- \rightarrow$ hadrons
4159 ± 20	BRANDELIK	78c DASP	$e^+ e^-$

<sup>1</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (293 \pm 57)^\circ$ .  
<sup>2</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects.  
<sup>3</sup> From a fit to Crystal Ball (OSTERHELD 86) data.  
<sup>4</sup> From a fit to BES (BAI 02c) data.

$\psi(4160)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>70 ± 10 OUR AVERAGE</b>			
65 +22 −16	AAIJ	13bc LHCB	$B^+ \rightarrow K^+ \mu^+ \mu^-$
71.8 ± 12.3	<sup>5</sup> ABLIKIM	08d BES2	$e^+ e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
79 ± 14	<sup>6</sup> MO	10 RVUE	$e^+ e^- \rightarrow$ hadrons
107 ± 10	<sup>7</sup> SETH	05A RVUE	$e^+ e^- \rightarrow$ hadrons
107 ± 16	<sup>8</sup> SETH	05A RVUE	$e^+ e^- \rightarrow$ hadrons
78 ± 20	BRANDELIK	78c DASP	$e^+ e^-$

<sup>5</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (293 \pm 57)^\circ$ .  
<sup>6</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects.  
<sup>7</sup> From a fit to Crystal Ball (OSTERHELD 86) data.  
<sup>8</sup> From a fit to BES (BAI 02c) data.

$\psi(4160)$  DECAY MODES

Due to the complexity of the  $c\bar{c}$  threshold region, in this listing, “seen” (“not seen”) means that a cross section for the mode in question has been measured at effective  $\sqrt{s}$  near this particle’s central mass value, more (less) than  $2\sigma$  above zero, without regard to any peaking behavior in  $\sqrt{s}$  or absence thereof. See mode listing(s) for details and references.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $e^+ e^-$	$(6.9 \pm 3.3) \times 10^{-6}$	
$\Gamma_2$ $\mu^+ \mu^-$	seen	
$\Gamma_3$ $D\bar{D}$	seen	
$\Gamma_4$ $D^0 \bar{D}^0$	seen	
$\Gamma_5$ $D^+ D^-$	seen	
$\Gamma_6$ $D^* \bar{D}^+ + \text{c.c.}$	seen	
$\Gamma_7$ $D^*(2007)^0 \bar{D}^0 + \text{c.c.}$	seen	
$\Gamma_8$ $D^*(2010)^+ D^- + \text{c.c.}$	seen	
$\Gamma_9$ $D^* \bar{D}^*$	seen	
$\Gamma_{10}$ $D^*(2007)^0 \bar{D}^*(2007)^0$	seen	
$\Gamma_{11}$ $D^*(2010)^+ D^*(2010)^-$	seen	
$\Gamma_{12}$ $D^0 D^- \pi^+ + \text{c.c. (excl. } D^*(2007)^0 \bar{D}^0 + \text{c.c., } D^*(2010)^+ D^- + \text{c.c.)}$	not seen	
$\Gamma_{13}$ $D \bar{D}^* \pi + \text{c.c. (excl. } D^* \bar{D}^*)$	seen	
$\Gamma_{14}$ $D^0 D^- \pi^+ + \text{c.c. (excl. } D^*(2010)^+ D^*(2010)^-)$	not seen	
$\Gamma_{15}$ $D_s^+ D_s^-$	not seen	
$\Gamma_{16}$ $D_s^{*+} \bar{D}_s^- + \text{c.c.}$	seen	
$\Gamma_{17}$ $J/\psi \pi^+ \pi^-$	$< 3$	$\times 10^{-3}$ 90%
$\Gamma_{18}$ $J/\psi \pi^0 \pi^0$	$< 3$	$\times 10^{-3}$ 90%
$\Gamma_{19}$ $J/\psi K^+ K^-$	$< 2$	$\times 10^{-3}$ 90%
$\Gamma_{20}$ $J/\psi \eta$	$< 8$	$\times 10^{-3}$ 90%
$\Gamma_{21}$ $J/\psi \pi^0$	$< 1$	$\times 10^{-3}$ 90%
$\Gamma_{22}$ $J/\psi \eta'$	$< 5$	$\times 10^{-3}$ 90%

$\Gamma_{23}$ $J/\psi \pi^+ \pi^- \pi^0$	$< 1$	$\times 10^{-3}$ 90%
$\Gamma_{24}$ $\psi(2S) \pi^+ \pi^-$	$< 4$	$\times 10^{-3}$ 90%
$\Gamma_{25}$ $\chi_{c1} \gamma$	$< 5$	$\times 10^{-3}$ 90%
$\Gamma_{26}$ $\chi_{c2} \gamma$	$< 1.3$	% 90%
$\Gamma_{27}$ $\chi_{c1} \pi^+ \pi^- \pi^0$	$< 2$	$\times 10^{-3}$ 90%
$\Gamma_{28}$ $\chi_{c2} \pi^+ \pi^- \pi^0$	$< 8$	$\times 10^{-3}$ 90%
$\Gamma_{29}$ $h_c(1P) \pi^+ \pi^-$	$< 5$	$\times 10^{-3}$ 90%
$\Gamma_{30}$ $h_c(1P) \pi^0 \pi^0$	$< 2$	$\times 10^{-3}$ 90%
$\Gamma_{31}$ $h_c(1P) \eta$	$< 2$	$\times 10^{-3}$ 90%
$\Gamma_{32}$ $h_c(1P) \pi^0$	$< 4$	$\times 10^{-4}$ 90%
$\Gamma_{33}$ $\phi \pi^+ \pi^-$	$< 2$	$\times 10^{-3}$ 90%
$\Gamma_{34}$ $\gamma X(3872) \rightarrow \gamma J/\psi \pi^+ \pi^-$	$< 6.8$	$\times 10^{-5}$ 90%
$\Gamma_{35}$ $\gamma X(3915) \rightarrow \gamma J/\psi \pi^+ \pi^-$	$< 1.36$	$\times 10^{-4}$ 90%
$\Gamma_{36}$ $\gamma X(3930) \rightarrow \gamma J/\psi \pi^+ \pi^-$	$< 1.18$	$\times 10^{-4}$ 90%
$\Gamma_{37}$ $\gamma X(3940) \rightarrow \gamma J/\psi \pi^+ \pi^-$	$< 1.47$	$\times 10^{-4}$ 90%
$\Gamma_{38}$ $\gamma X(3872) \rightarrow \gamma \gamma J/\psi$	$< 1.05$	$\times 10^{-4}$ 90%
$\Gamma_{39}$ $\gamma X(3915) \rightarrow \gamma \gamma J/\psi$	$< 1.26$	$\times 10^{-4}$ 90%
$\Gamma_{40}$ $\gamma X(3930) \rightarrow \gamma \gamma J/\psi$	$< 8.8$	$\times 10^{-5}$ 90%
$\Gamma_{41}$ $\gamma X(3940) \rightarrow \gamma \gamma J/\psi$	$< 1.79$	$\times 10^{-4}$ 90%
$\Gamma_{42}$ $K^+ K^-$		

$\psi(4160)$  PARTIAL WIDTHS

$\Gamma(e^+ e^-)$	DOCUMENT ID	TECN	COMMENT
VALUE (keV)			
<b>0.48 ± 0.22</b>	<sup>9</sup> ABLIKIM	08d BES2	$e^+ e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.4 to 1.1	<sup>10</sup> MO	10 RVUE	$e^+ e^- \rightarrow$ hadrons
0.83 ± 0.08	<sup>11</sup> SETH	05A RVUE	$e^+ e^- \rightarrow$ hadrons
0.84 ± 0.13	<sup>12</sup> SETH	05A RVUE	$e^+ e^- \rightarrow$ hadrons
0.77 ± 0.23	BRANDELIK	78c DASP	$e^+ e^-$

<sup>9</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (293 \pm 57)^\circ$ .  
<sup>10</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects. Four sets of solutions are obtained with the same fit quality, mass and total width, but with different  $e^+ e^-$  partial widths. We quote only the range of values.  
<sup>11</sup> From a fit to Crystal Ball (OSTERHELD 86) data.  
<sup>12</sup> From a fit to BES (BAI 02c) data.

$\psi(4160)$   $\Gamma(i) \times \Gamma(e^+ e^-)/\Gamma(\text{total})$

$\Gamma(\chi_{c1} \gamma) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (eV)	CL%		
<b>&lt; 2.2</b>	90	<sup>13</sup> HAN	15 BELL $10.58 e^+ e^- \rightarrow \chi_{c1} \gamma$
<sup>13</sup> Using $B(\eta \rightarrow \gamma \gamma) = (39.41 \pm 0.21)\%$ .			

$\Gamma(\chi_{c2} \gamma) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (eV)	CL%		
<b>&lt; 6.1</b>	90	<sup>14</sup> HAN	15 BELL $10.58 e^+ e^- \rightarrow \chi_{c2} \gamma$
<sup>14</sup> Using $B(\eta \rightarrow \gamma \gamma) = (39.41 \pm 0.21)\%$ .			

$\psi(4160)$   $\Gamma(i) \times \Gamma(e^+ e^-)/\Gamma^2(\text{total})$

$\Gamma(J/\psi \eta)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (units $10^{-8}$ )			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$2.8 \pm 0.9 \pm 0.9$	<sup>15</sup> WANG	13B BELL	$e^+ e^- \rightarrow J/\psi \eta \gamma$
$12.8 \pm 1.7 \pm 2.0$	<sup>16</sup> WANG	13B BELL	$e^+ e^- \rightarrow J/\psi \eta \gamma$
<sup>15</sup> Solution I of two equivalent solutions in a fit using two interfering resonances. Mass and width fixed at 4153 MeV and 103 MeV, respectively. <sup>16</sup> Solution II of two equivalent solutions in a fit using two interfering resonances. Mass and width fixed at 4153 MeV and 103 MeV, respectively.			

$\psi(4160)$  BRANCHING RATIOS

$\Gamma(\mu^+ \mu^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE			
<b>seen</b>	<sup>17</sup> AAIJ	13bc LHCB	$B^+ \rightarrow K^+ \mu^+ \mu^-$
<sup>17</sup> AAIJ 13bc report $B(B^+ \rightarrow K^+ \psi(4160)) B(\psi(4160) \rightarrow \mu^+ \mu^-) = (3.5^{+0.9}_{-0.8}) \times 10^{-9}$ .			
$\Gamma(D\bar{D})/\Gamma(D^* \bar{D}^*)$	DOCUMENT ID	TECN	COMMENT
VALUE			
<b>0.02 ± 0.03 ± 0.02</b>	AUBERT	09M BABR	$e^+ e^- \rightarrow \gamma D^{(*)} \bar{D}^{(*)}$



See key on page 601

Meson Particle Listings

$\psi(4160)$ ,  $X(4160)$ ,  $X(4200)^\pm$ ,  $X(4230)$

$\Gamma(\gamma X(3915) \rightarrow \gamma J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{35}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<1.36 \times 10^{-4}$	90	24 XIAO	13 $\psi(4160) \rightarrow \gamma J/\psi \pi^+ \pi^-$
24 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(\gamma X(3930) \rightarrow \gamma J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{36}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<1.18 \times 10^{-4}$	90	25 XIAO	13 $\psi(4160) \rightarrow \gamma J/\psi \pi^+ \pi^-$
25 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(\gamma X(3940) \rightarrow \gamma J/\psi \pi^+ \pi^-)/\Gamma_{\text{total}}$   $\Gamma_{37}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<1.47 \times 10^{-4}$	90	26 XIAO	13 $\psi(4160) \rightarrow \gamma J/\psi \pi^+ \pi^-$
26 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(\gamma X(3872) \rightarrow \gamma \gamma J/\psi)/\Gamma_{\text{total}}$   $\Gamma_{38}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<1.05 \times 10^{-4}$	90	27 XIAO	13 $\psi(4160) \rightarrow \gamma \gamma J/\psi$
27 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(\gamma X(3915) \rightarrow \gamma \gamma J/\psi)/\Gamma_{\text{total}}$   $\Gamma_{39}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<1.26 \times 10^{-4}$	90	28 XIAO	13 $\psi(4160) \rightarrow \gamma \gamma J/\psi$
28 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(\gamma X(3930) \rightarrow \gamma \gamma J/\psi)/\Gamma_{\text{total}}$   $\Gamma_{40}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<0.88 \times 10^{-4}$	90	29 XIAO	13 $\psi(4160) \rightarrow \gamma \gamma J/\psi$
29 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(\gamma X(3940) \rightarrow \gamma \gamma J/\psi)/\Gamma_{\text{total}}$   $\Gamma_{41}/\Gamma$

VALUE	CL%	DOCUMENT ID	COMMENT
$<1.79 \times 10^{-4}$	90	30 XIAO	13 $\psi(4160) \rightarrow \gamma \gamma J/\psi$
30 Obtained by analyzing CLEO data but not authored by the CLEO Collaboration.			

$\Gamma(K^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{42}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2 \times 10^{-5}$	90	31 DRUZHININ	15 RVUE	$e^+ e^- \rightarrow \psi(3770)$
31 DRUZHININ 15 uses BABAR and CLEO data takitaking into account interference of the processes $e^+ e^- \rightarrow K^+ K^-$ and $e^+ e^- \rightarrow K_S^0 K_L^0$ .				

$\psi(4160)$  REFERENCES

ABLIKIM	15L	PR D91 112005	M. Ablikim <i>et al.</i>	(BES III Collab.)
DRUZHININ	15	PR D92 054024	V.P. Druzhinin	(NOVO)
HAN	15	PR D92 012011	Y.L. Han <i>et al.</i>	(BELLE Collab.)
AAIJ	13BC	PRL 111 112003	R. Aaij <i>et al.</i>	(LHCb Collab.)
WANG	13B	PR D87 051101	X.L. Wang <i>et al.</i>	(BELLE Collab.)
XIAO	13	PR D87 057501	T. Xiao <i>et al.</i>	(NWES, WAYN)
PAKHLOVA	11	PR D83 011101	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PEDLAR	11	PRL 107 041803	T. Pedlar <i>et al.</i>	(CLEO Collab.)
DEL-AMO-SA...	10N	PR D82 052004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
MO	10	PR D82 077501	X.H. Mo, C.Z. Yuan, P. Wang	(BABAR) (BHEP)
AUBERT	09M	PR D79 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)
CRONIN-HEN...	09	PR D80 072001	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
PAKHLOVA	09	PR D80 091101	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
ABLIKIM	08D	PL B660 315	M. Ablikim <i>et al.</i>	(BES Collab.)
PAKHLOVA	08	PR D77 011103	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PAKHLOVA	08A	PRL 100 062001	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PAKHLOVA	07	PRL 98 092001	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
COAN	06	PRL 96 162003	T.E. Coan <i>et al.</i>	(CLEO Collab.)
SETH	05A	PR D72 017501	K.K. Seth	(BES Collab.)
BAI	02C	PRL 88 101802	J.Z. Bai <i>et al.</i>	(BES Collab.)
BAI	00	PRL 84 594	J.Z. Bai <i>et al.</i>	(BES Collab.)
OSTERHELD	86	SLAC-PUB-4160	A. Osterheld <i>et al.</i>	(SLAC Crystal Ball Collab.)
BRANDELIK	78C	PL 76B 361	R. Brandelik <i>et al.</i>	(DASP Collab.)

$X(4160)$

$I^G(J^{PC}) = ?^?(?^{??})$

OMITTED FROM SUMMARY TABLE

Seen by PAKHLOV 08 in  $e^+ e^- \rightarrow J/\psi X$ ,  $X \rightarrow D^* \bar{D}^*$

$X(4160)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$4156^{+25}_{-20} \pm 15$	24	PAKHLOV	08 BELL	$e^+ e^- \rightarrow J/\psi X$

$X(4160)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$139^{+111}_{-61} \pm 21$	24	PAKHLOV	08 BELL	$e^+ e^- \rightarrow J/\psi X$

$X(4160)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $D \bar{D}$	not seen
$\Gamma_2$ $D^* \bar{D} + \text{c.c.}$	not seen
$\Gamma_3$ $D^* \bar{D}^*$	seen

$X(4160)$  BRANCHING RATIOS

$\Gamma(D\bar{D})/\Gamma(D^*\bar{D}^*)$				$\Gamma_1/\Gamma_3$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.09</b>	90	PAKHLOV	08 BELL	$e^+e^- \rightarrow J/\psi X$

$\Gamma(D^*\bar{D} + \text{c.c.})/\Gamma(D^*\bar{D}^*)$				$\Gamma_2/\Gamma_3$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.22</b>	90	PAKHLOV	08 BELL	$e^+e^- \rightarrow J/\psi X$

$X(4160)$  REFERENCES

PAKHLOV	08	PRL 100 202001	P. Pakhlov <i>et al.</i>	(BELLE Collab.)
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$X(4200)^\pm$

$I(J^P) = ?(1^+)$

OMITTED FROM SUMMARY TABLE

Reported by CHILIKIN 14 in  $J/\psi \pi^+$  at a significance of  $6.2\sigma$ . Assignments of  $0^-$ ,  $1^-$ ,  $2^-$ , and  $2^+$  excluded at  $6.1\sigma$ ,  $7.4\sigma$ ,  $4.4\sigma$ , and  $7.0\sigma$  level, respectively. Needs confirmation.

$X(4200)^\pm$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$4196^{+31+17}_{-29-13}$	CHILIKIN	14 BELL	$\bar{B}^0 \rightarrow J/\psi K^- \pi^+$

$X(4200)^\pm$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$370 \pm 70^{+70}_{-132}$	CHILIKIN	14 BELL	$\bar{B}^0 \rightarrow J/\psi K^- \pi^+$

$X(4200)^\pm$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $J/\psi \pi^+$	seen

$X(4200)^\pm$  BRANCHING RATIOS

$\Gamma(J/\psi\pi^+)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	CHILIKIN	14 BELL	$\bar{B}^0 \rightarrow J/\psi K^- \pi^+$

$X(4200)^\pm$  REFERENCES

CHILIKIN	14	PR D90 112009	K. Chilikin <i>et al.</i>	(BELLE Collab.)
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$X(4230)$

$I^G(J^{PC}) = ?^?(1^- -)$

OMITTED FROM SUMMARY TABLE

Enhancement reported by ABLIKIM 15C in  $e^+ e^- \rightarrow \omega \chi_{c0}$  at  $\sqrt{s} = 4.23\text{--}4.26$  GeV at  $9\sigma$  significance. Lineshape found to be inconsistent with origination from  $X(4260)$ . NEEDS CONFIRMATION.

$X(4230)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$4230 \pm 8 \pm 6$	180	1 ABLIKIM	15c BES3	$e^+ e^- \rightarrow \omega \chi_{c0}$

1 From a 3-parameter fit of measured cross sections from  $\sqrt{s} = 4.21\text{--}4.42$  GeV to a phase-space modified Breit-Wigner function, using the decays  $\chi_{c0} \rightarrow \pi^+ \pi^-$ ,  $\chi_{c0} \rightarrow K^+ K^-$ , and  $\omega \rightarrow \pi^+ \pi^- \pi^0$ .

$X(4230)$  00WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$38 \pm 12 \pm 2$	180	1 ABLIKIM	15c BES3	$e^+ e^- \rightarrow \omega \chi_{c0}$

1 From a 3-parameter fit of measured cross sections from  $\sqrt{s} = 4.21\text{--}4.42$  GeV to a phase-space modified Breit-Wigner function, using the decays  $\chi_{c0} \rightarrow \pi^+ \pi^-$ ,  $\chi_{c0} \rightarrow K^+ K^-$ , and  $\omega \rightarrow \pi^+ \pi^- \pi^0$ .

Meson Particle Listings

X(4230), X(4240)<sup>±</sup>, X(4250)<sup>±</sup>, X(4260)

X(4230) DECAY MODES				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$e^+ e^-$	seen		
$\Gamma_2$	$\omega \chi_{c0}$			
X(4230) $\Gamma(i)\Gamma(e^+ e^-)/\Gamma(\text{total})$				
$\Gamma(\omega \chi_{c0}) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$		$\Gamma_2 \Gamma_1/\Gamma$		
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
$2.7 \pm 0.5 \pm 0.4$	180	<sup>1</sup> ABLIKIM	15c BES3	$e^+ e^- \rightarrow \omega \chi_{c0}$
<sup>1</sup> From a 3-parameter fit of measured cross sections from $\sqrt{s} = 4.21\text{--}4.42$ GeV to a phase-space modified Breit-Wigner function, using the decays $\chi_{c0} \rightarrow \pi^+ \pi^-$ , $\chi_{c0} \rightarrow K^+ K^-$ , and $\omega \rightarrow \pi^+ \pi^- \pi^0$ .				

X(4230) BRANCHING RATIOS				
Γ(ω χ <sub>c0</sub> )/Γ <sub>total</sub>		Γ <sub>2</sub> /Γ		
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	180	<sup>1</sup> ABLIKIM	15c BES3	e <sup>+</sup> e <sup>-</sup> → ω χ <sub>c0</sub>
<sup>1</sup> From a 3-parameter fit of measured cross sections from √s = 4.21–4.42 GeV to a phase-space modified Breit-Wigner function, using the decays χ <sub>c0</sub> → π <sup>+</sup> π <sup>-</sup> , χ <sub>c0</sub> → K <sup>+</sup> K <sup>-</sup> , and ω → π <sup>+</sup> π <sup>-</sup> π <sup>0</sup> .				

X(4230) REFERENCES				
ABLIKIM	15C	PRL 114 092003	M. Ablikim <i>et al.</i>	(BES III Collab.)

X(4240)<sup>±</sup>

$I^G(J^P) = ?^?(0^-)$

OMITTED FROM SUMMARY TABLE

Spin and parity assignment  $J^P = 0^-$  is favored over  $1^-$ ,  $2^-$ , and  $2^+$  by 8 σ and over  $1^+$  by 1 σ, according to the four-dimensional amplitude analysis of AAIJ 14AG.

X(4240) <sup>±</sup> MASS						
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT			
4239 ± 18 <sup>+45</sup> <sub>-10</sub>	<sup>1</sup> AAIJ	14AG LHCB	$B^0 \rightarrow K^+ \pi^- \psi(2S)$			
<sup>1</sup> From a 4-dimensional analysis when a second, lower mass resonance is allowed in the X(4430) <sup>±</sup> fit, with significance 6 $\sigma$ including systematic variations.						
X(4240) <sup>±</sup> WIDTH						
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT			
220 ± 47 <sup>+108</sup> <sub>-74</sub>	<sup>2</sup> AAIJ	14AG LHCB	$B^0 \rightarrow K^+ \pi^- \psi(2S)$			
<sup>2</sup> From a 4-dimensional analysis when a second, lower mass resonance is allowed in the X(4430) <sup>±</sup> fit, with significance 6 $\sigma$ including systematic variations.						

X(4240) <sup>±</sup> DECAY MODES				
Mode		Fraction (Γ <sub>i</sub> /Γ)		
Γ <sub>1</sub>	π <sup>−</sup> ψ(2S)	seen		

X(4240) <sup>±</sup> BRANCHING RATIOS				
Γ(π <sup>−</sup> ψ(2S))/Γ <sub>total</sub>		Γ <sub>1</sub> /Γ		
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	<sup>3</sup> AAIJ	14AG LHCB	B <sup>0</sup> → K <sup>+</sup> π <sup>−</sup> ψ(2S)	
<sup>3</sup> From a 4-dimensional analysis when a second, lower mass resonance is allowed in the X(4430) <sup>±</sup> fit. No partial branching fraction quoted.				

X(4240) <sup>±</sup> REFERENCES				
AAIJ	14AG	PRL 112 222002	R. Aaij <i>et al.</i>	(LHCb Collab.)

X(4250)<sup>±</sup>

$I(J^P) = ?(?^?)$

OMITTED FROM SUMMARY TABLE

Observed by MIZUK 08 in the π<sup>+</sup> χ<sub>c1</sub>(1P) invariant mass distribution in  $\bar{B}^0 \rightarrow K^- \pi^+ \chi_{c1}(1P)$  decays. Not seen by LEES 12B in this same mode after accounting for K π resonant mass and angular structure.

X(4250) <sup>±</sup> MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
4248 +44 +180 -29 -35	<sup>1</sup> MIZUK	08 BELL	$\bar{B}^0 \rightarrow K^- \pi^+ \chi_{c1}(1P)$	
<sup>1</sup> From a Dalitz plot analysis with two Breit-Wigner amplitudes.				

X(4250) <sup>±</sup> WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
177 <sup>+54+316</sup> <sub>-39-61</sub>	<sup>2</sup> MIZUK	08 BELL	$\overline{B}^0 \rightarrow K^- \pi^+ \chi_{c1}(1P)$	
<sup>2</sup> From a Dalitz plot analysis with two Breit-Wigner amplitudes.				

X(4250) <sup>±</sup> DECAY MODES				
Mode		Fraction (Γ <sub>i</sub> /Γ)		
Γ <sub>1</sub>	π <sup>+</sup> χ <sub>c1</sub> (1P)	seen		

X(4250) <sup>±</sup> BRANCHING RATIOS				
Γ(π <sup>+</sup> χ <sub>c1</sub> (1P))/Γ <sub>total</sub>			Γ <sub>1</sub> /Γ	
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	<sup>3</sup> MIZUK	08 BELL	$\bar{B}^0 \rightarrow K^- \pi^+ \chi_{c1}(1P)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen	<sup>4</sup> LEES	12B BABR	$B \rightarrow K \pi \chi_{c1}(1P)$	
<sup>3</sup> With a product branching fraction measurement of B( $\bar{B}^0 \rightarrow K^- X(4250)^+$ ) × B(X(4250) <sup>+</sup> → π <sup>+</sup> χ <sub>c1</sub> (1P)) = (4.0 <sup>+2.3</sup> <sub>-0.9</sub> <sup>+19.7</sup> <sub>-0.5</sub> ) × 10 <sup>-5</sup> .				
<sup>4</sup> With a product branching fraction limit of B( $\bar{B}^0 \rightarrow X(4250)^+ K^-$ ) × B(X(4250) <sup>+</sup> → χ <sub>c1</sub> π <sup>+</sup> ) < 4.0 × 10 <sup>-5</sup> at 90% CL.				

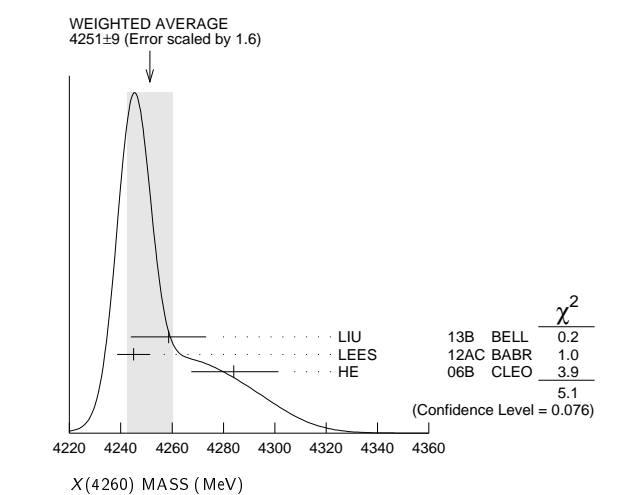
X(4250) <sup>±</sup> REFERENCES				
LEES MIZUK	12B 08	PR D85 052003 PR D78 072004	J.P. Lees <i>et al.</i> R. Mizuk <i>et al.</i>	(BABAR Collab.) (BELLE Collab.)

X(4260)

$I^G(J^{PC}) = ?^?(1^- -)$

Seen in radiative return from e<sup>+</sup> e<sup>-</sup> collisions at √s = 9.54–10.58 GeV by AUBERT,B 05l, HE 06B, and YUAN 07, and in e<sup>+</sup> e<sup>-</sup> collisions at √s ≈ 4.26 GeV by COAN 06. Possibly seen by AUBERT 06 in  $B^- \rightarrow K^- \pi^+ \pi^- J/\psi$ . See also the mini-review under the X(3872). (See the index for the page number.)

X(4260) MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
4251 ± 9	OUR AVERAGE	Error includes scale factor of 1.6. See the ideogram below.		
4258.6 ± 8.3 ± 12.1	<sup>1</sup> LIU	13B BELL	e <sup>+</sup> e <sup>-</sup> → γ π <sup>+</sup> π <sup>-</sup> J/ψ	
4245 ± 5 ± 4	<sup>2</sup> LEES	12AC BABR	10.58 e <sup>+</sup> e <sup>-</sup> → γ π <sup>+</sup> π <sup>-</sup> J/ψ	
4284 <sup>+17</sup> <sub>-16</sub> ± 413.6	HE	06B CLEO	9.4-10.6 e <sup>+</sup> e <sup>-</sup> → γ π <sup>+</sup> π <sup>-</sup> J/ψ	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
4247 ± 12 <sup>+17</sup> <sub>-32</sub>	<sup>1,3</sup> YUAN	07 BELL	10.58 e <sup>+</sup> e <sup>-</sup> → γ π <sup>+</sup> π <sup>-</sup> J/ψ	
4259 ± 8 ± <sup>2</sup> <sub>6</sub> 125	<sup>4</sup> AUBERT,B	05l BABR	10.58 e <sup>+</sup> e <sup>-</sup> → γ π <sup>+</sup> π <sup>-</sup> J/ψ	



# Meson Particle Listings

## X(4260)

- <sup>1</sup> From a two-resonance fit.  
<sup>2</sup> From a single-resonance fit. Supersedes AUBERT,B 05i.  
<sup>3</sup> Superseded by LIU 13B.  
<sup>4</sup> From a single-resonance fit. Two interfering resonances are not excluded. Superseded by LEES 12Ac.

### X(4260) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>120 ± 12 OUR AVERAGE</b>				Error includes scale factor of 1.1.
134.1 ± 16.4 ± 5.5	<sup>1</sup> LIU	13B BELL		$e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$
114 $^{+16}_{-15} \pm 7$	<sup>2</sup> LEES	12Ac BABR	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
73 $^{+39}_{-25} \pm 5$ 13.6	HE	06B CLEO	9.4–10.6 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
108 ± 19 ± 10	<sup>1,3</sup> YUAN	07 BELL	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
88 ± 23 $^{+6}_{-4}$ 125	<sup>4</sup> AUBERT,B	05i BABR	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
<sup>1</sup> From a two-resonance fit. <sup>2</sup> From a single-resonance fit. Supersedes AUBERT,B 05i. <sup>3</sup> Superseded by LIU 13B. <sup>4</sup> From a single-resonance fit. Two interfering resonances are not excluded. Superseded by LEES 12Ac.				

### X(4260) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 e^+e^-$	
$\Gamma_2 J/\psi\pi^+\pi^-$	seen
$\Gamma_3 J/\psi f_0(980), f_0(980) \rightarrow \pi^+\pi^-$	seen
$\Gamma_4 X(3900)^\pm\pi^\mp, X^\pm \rightarrow J/\psi\pi^\pm$	seen
$\Gamma_5 J/\psi\pi^0\pi^0$	seen
$\Gamma_6 J/\psi K^+K^-$	seen
$\Gamma_7 J/\psi K_S^0 K_S^0$	not seen
$\Gamma_8 X(3872)\gamma$	seen
$\Gamma_9 J/\psi\eta$	not seen
$\Gamma_{10} J/\psi\pi^0$	not seen
$\Gamma_{11} J/\psi\eta'$	not seen
$\Gamma_{12} J/\psi\pi^+\pi^-\pi^0$	not seen
$\Gamma_{13} J/\psi\eta\pi^0$	not seen
$\Gamma_{14} J/\psi\eta\eta$	not seen
$\Gamma_{15} \psi(2S)\pi^+\pi^-$	not seen
$\Gamma_{16} \psi(2S)\eta$	not seen
$\Gamma_{17} \chi_{c0}\omega$	not seen
$\Gamma_{18} \chi_{c1}\gamma$	not seen
$\Gamma_{19} \chi_{c2}\gamma$	not seen
$\Gamma_{20} \chi_{c1}\pi^+\pi^-\pi^0$	not seen
$\Gamma_{21} \chi_{c2}\pi^+\pi^-\pi^0$	not seen
$\Gamma_{22} h_c(1P)\pi^+\pi^-$	not seen
$\Gamma_{23} \phi\pi^+\pi^-$	not seen
$\Gamma_{24} \phi f_0(980) \rightarrow \phi\pi^+\pi^-$	not seen
$\Gamma_{25} D\bar{D}$	not seen
$\Gamma_{26} D^0\bar{D}^0$	not seen
$\Gamma_{27} D^+\bar{D}^-$	not seen
$\Gamma_{28} D^*\bar{D}^+ + c.c.$	not seen
$\Gamma_{29} D^*(2007)^0\bar{D}^0 + c.c.$	not seen
$\Gamma_{30} D^*(2010)^+ D^- + c.c.$	not seen
$\Gamma_{31} D^*\bar{D}^*$	not seen
$\Gamma_{32} D^*(2007)^0\bar{D}^*(2007)^0$	not seen
$\Gamma_{33} D^*(2010)^+ D^*(2010)^-$	not seen
$\Gamma_{34} D\bar{D}\pi + c.c.$	
$\Gamma_{35} D^0 D^- \pi^+ + c.c. \text{ (excl. } D^*(2007)^0\bar{D}^{*0} + c.c., D^*(2010)^+ D^- + c.c.)$	not seen
$\Gamma_{36} D\bar{D}^* \pi + c.c. \text{ (excl. } D^*\bar{D}^*)$	not seen
$\Gamma_{37} D^0 D^{*-} \pi^+ + c.c. \text{ (excl. } D^*(2010)^+ D^*(2010)^-)$	not seen
$\Gamma_{38} D^0 D^*(2010)^- \pi^+ + c.c.$	not seen
$\Gamma_{39} D^*\bar{D}^* \pi$	not seen
$\Gamma_{40} D_s^+ D_s^-$	not seen
$\Gamma_{41} D_s^{*+} D_s^- + c.c.$	not seen
$\Gamma_{42} D_s^{*+} D_s^{*-}$	not seen
$\Gamma_{43} \rho\bar{\rho}$	not seen
$\Gamma_{44} K_S^0 K^\pm \pi^\mp$	not seen
$\Gamma_{45} K^+ K^- \pi^0$	not seen

### X(4260) $\Gamma(i) \times \Gamma(e^+e^-)/\Gamma(\text{total})$

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.2 ± 1.0 OUR AVERAGE</b>				
9.2 ± 0.8 ± 0.7	<sup>1</sup> LEES	12Ac BABR	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
8.9 $^{+3.9}_{-3.1} \pm 1.8$	8.1 HE	06B CLEO	9.4–10.6 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.4 ± 0.8 ± 0.6	<sup>2</sup> LIU	13B BELL	$e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
20.5 ± 1.4 ± 2.0	<sup>3</sup> LIU	13B BELL	$e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
6.0 ± 1.2 $^{+4.7}_{-0.5}$	<sup>2,4</sup> YUAN	07 BELL	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
20.6 ± 2.3 $^{+9.1}_{-1.7}$	<sup>3,4</sup> YUAN	07 BELL	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
5.5 ± 1.0 $^{+0.8}_{-0.7}$	125 <sup>5</sup> AUBERT,B	05i BABR	10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^- J/\psi$	
<sup>1</sup> From a single-resonance fit. Supersedes AUBERT,B 05i. <sup>2</sup> Solution I of two equivalent solutions in a fit using two interfering resonances. <sup>3</sup> Solution II of two equivalent solutions in a fit using two interfering resonances. <sup>4</sup> Superseded by LIU 13B. <sup>5</sup> From a single-resonance fit. Two interfering resonances are not excluded. Superseded by LEES 12Ac.				

### $\Gamma(J/\psi K^+ K^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.7</b>	90	<sup>1</sup> SHEN	14 BELL	9.4–10.9 $e^+e^- \rightarrow \gamma K^+ K^- J/\psi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.2	90	<sup>2</sup> YUAN	08 BELL	$e^+e^- \rightarrow \gamma K^+ K^- J/\psi$
<sup>1</sup> From a fit of the broad $K^+ K^- J/\psi$ enhancement including a coherent X(4260) amplitude with mass and width from LIU 13B. Supersedes YUAN 08. <sup>2</sup> From a fit of the broad $K^+ K^- J/\psi$ enhancement including a coherent X(4260) amplitude with mass and width from YUAN 07.				

### $\Gamma(J/\psi K_S^0 K_S^0) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.85</b>	90	<sup>1</sup> SHEN	14 BELL	9.4–10.9 $e^+e^- \rightarrow \gamma K_S^0 K_S^0 J/\psi$
<sup>1</sup> From a fit of the $K_S^0 K_S^0 J/\psi$ mass range from 4.4 to 5.5 GeV including a coherent X(4260) amplitude with mass and width from LIU 13B.				

### $\Gamma(J/\psi\eta) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<14.2	90	WANG	13B BELL	$e^+e^- \rightarrow J/\psi\eta\eta$

### $\Gamma(\psi(2S)\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<4.3	90	<sup>1</sup> LIU	08H RVUE	10.58 $e^+e^- \rightarrow \psi(2S)\pi^+\pi^-\gamma$
7.4 $^{+2.1}_{-1.7}$		<sup>2</sup> LIU	08H RVUE	10.58 $e^+e^- \rightarrow \psi(2S)\pi^+\pi^-\gamma$
<sup>1</sup> For constructive interference with the X(4360) in a combined fit of AUBERT 07s and WANG 07b data with three resonances. <sup>2</sup> For destructive interference with the X(4360) in a combined fit of AUBERT 07s and WANG 07b data with three resonances.				

### $\Gamma(\chi_{c1}\gamma) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.4</b>	90	<sup>1</sup> HAN	15 BELL	10.58 $e^+e^- \rightarrow \chi_{c1}\gamma$
<sup>1</sup> Using $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.21)\%$ .				

### $\Gamma(\chi_{c2}\gamma) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.0</b>	90	<sup>1</sup> HAN	15 BELL	10.58 $e^+e^- \rightarrow \chi_{c2}\gamma$
<sup>1</sup> Using $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.21)\%$ .				

### $\Gamma(\phi\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.4</b>	90	AUBERT,BE	06D BABR	10.6 $e^+e^- \rightarrow K^+ K^- \pi^+ \pi^- \gamma$

### $\Gamma(\phi f_0(980) \rightarrow \phi\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.29</b>	90	<sup>1</sup> AUBERT	07AK BABR	10.6 $e^+e^- \rightarrow \pi^+\pi^- K^+ K^- \gamma$
<sup>1</sup> AUBERT 07AK reports $[\Gamma(X(4260) \rightarrow \phi f_0(980) \rightarrow \phi\pi^+\pi^-) \times \Gamma(X(4260) \rightarrow e^+e^-)/\Gamma_{\text{total}}] \times [B(\phi(1020) \rightarrow K^+ K^-)] < 0.14$ eV which we divide by our best value $B(\phi(1020) \rightarrow K^+ K^-) = 48.9 \times 10^{-2}$ .				

### $\Gamma(K_S^0 K^\pm \pi^\mp) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.5	90	AUBERT	08s BABR	10.6 $e^+e^- \rightarrow K_S^0 K^\pm \pi^\mp \gamma$



See key on page 601

Meson Particle Listings  
X(4260), X(4350), X(4360)

$\Gamma(D_s^{*+} D_s^{*-})/\Gamma_{\text{total}}$		$\Gamma_{42}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
not seen	CRONIN-HEN..09	CLEO	$e^+ e^- \rightarrow D_s^{*+} D_s^{*-}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	PAKHLOVA 11	BELL	$e^+ e^- \rightarrow D_s^{*+} D_s^{*-} \gamma$
not seen	DEL-AMO-SA..10N	BABR	$e^+ e^- \rightarrow D_s^{*+} D_s^{*-} \gamma$
$\Gamma(D_s^{*+} D_s^{*-})/\Gamma(J/\psi \pi^+ \pi^-)$		$\Gamma_{42}/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
< 9.5	90	CRONIN-HEN..09	CLEO $e^+ e^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<30	95	DEL-AMO-SA..10N	BABR $10.6 e^+ e^-$
$\Gamma(p\overline{p})/\Gamma(J/\psi \pi^+ \pi^-)$		$\Gamma_{43}/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	COMMENT
<0.13	90	<sup>1</sup> AUBERT 06B	$e^+ e^- \rightarrow p\overline{p} \gamma$
<sup>1</sup> Using 4259 ± 10 MeV for the mass and 88 ± 24 MeV for the width of X(4260).			

X(4260) REFERENCES

ABLIKIM 15Q	PR D92 012008	M. Ablikim <i>et al.</i>	(BES III Collab.)
HAN 15	PR D92 012011	Y.L. Han <i>et al.</i>	(BELLE Collab.)
ABLIKIM 14	PRL 112 092001	M. Ablikim <i>et al.</i>	(BES III Collab.)
SHEN 14	PR D89 072015	C.P. Shen <i>et al.</i>	(BELLE Collab.)
ABLIKIM 13T	PRL 110 252001	M. Ablikim <i>et al.</i>	(BES III Collab.)
LIU 13B	PRL 110 252002	Z.Q. Liu <i>et al.</i>	(BELLE Collab.)
WANG 13B	PR D87 051101	X.L. Wang <i>et al.</i>	(BELLE Collab.)
LEES 12AC	PR D86 051102	J.P. Lees <i>et al.</i>	(BABAR Collab.)
PAKHLOVA 11	PR D83 011101	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PEDLAR 11	PRL 107 041803	T. Pedlar <i>et al.</i>	(CLEO Collab.)
DEL-AMO-SA..10N	PR D52 052004	P. del Amo Sanchez <i>et al.</i>	(BABAR Collab.)
AUBERT 09M	PR D79 092001	B. Aubert <i>et al.</i>	(BABAR Collab.)
CRONIN-HEN...09	PR D80 072001	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
PAKHLOVA 09	PR D80 091101	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
AUBERT 08S	PR D77 092002	B. Aubert <i>et al.</i>	(BABAR Collab.)
LIU 08H	PR D78 014032	Z.Q. Liu, X.S. Qin, C.Z. Yuan	
PAKHLOVA 08	PR D77 011103	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
PAKHLOVA 08A	PRL 100 062001	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
YUAN 08	PR D77 011105	C.Z. Yuan <i>et al.</i>	(BELLE Collab.)
AUBERT 07AK	PR D76 012008	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT 07BE	PR D76 111105	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT 07S	PRL 98 212001	B. Aubert <i>et al.</i>	(BABAR Collab.)
PAKHLOVA 07	PRL 98 092001	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
WANG 07D	PRL 99 142002	X.L. Wang <i>et al.</i>	(BELLE Collab.)
YUAN 07	PRL 99 182004	C.Z. Yuan <i>et al.</i>	(BELLE Collab.)
AUBERT 06	PR D73 011101	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT 06B	PR D73 012005	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,BE 06D	PR D74 091103	B. Aubert <i>et al.</i>	(BABAR Collab.)
COAN 06	PRL 96 162003	T.E. Coan <i>et al.</i>	(CLEO Collab.)
HE 06B	PR D74 091104	Q. He <i>et al.</i>	(CLEO Collab.)
AUBERT,B 05I	PRL 95 142001	B. Aubert <i>et al.</i>	(BABAR Collab.)

X(4350)

$I^G(J^{PC}) = 0^+(?^{?+})$

OMITTED FROM SUMMARY TABLE  
Seen by SHEN 10 in the  $\gamma\gamma \rightarrow J/\psi\phi$ . Needs confirmation.

X(4350) MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4350.6<sup>+4.6</sup><sub>-5.1</sub> ± 0.7</b>	8.8 <sup>+4.2</sup> <sub>-3.2</sub>	<sup>1</sup> SHEN 10	BELL	10.6 $e^+ e^- \rightarrow e^+ e^- J/\psi\phi$
<sup>1</sup> Statistical significance of 3.2 $\sigma$ .				

X(4350) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>13<sup>+18</sup><sub>-9</sub> ± 4</b>	8.8 <sup>+4.2</sup> <sub>-3.2</sub>	<sup>1</sup> SHEN 10	BELL	10.6 $e^+ e^- \rightarrow e^+ e^- J/\psi\phi$
<sup>1</sup> Statistical significance of 3.2 $\sigma$ .				

X(4350) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $J/\psi\phi$	seen
$\Gamma_2$ $\gamma\gamma$	seen

X(4350)  $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(\gamma\gamma) \times \Gamma(J/\psi\phi)/\Gamma_{\text{total}}$	VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2\Gamma_1/\Gamma$
<b>6.7<sup>+3.2</sup><sub>-2.4</sub> ± 1.1</b>	8.8 <sup>+4.2</sup> <sub>-3.2</sub>	<sup>1</sup> SHEN 10	BELL	10.6 $e^+ e^- \rightarrow e^+ e^- J/\psi\phi$		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
1.5 <sup>+0.7</sup> <sub>-0.6</sub> ± 0.3	8.8 <sup>+4.2</sup> <sub>-3.2</sub>	<sup>2</sup> SHEN 10	BELL	10.6 $e^+ e^- \rightarrow e^+ e^- J/\psi\phi$		
<sup>1</sup> For $J^P = 0^+$ . Statistical significance of 3.2 $\sigma$ . <sup>2</sup> For $J^P = 2^+$ . Statistical significance of 3.2 $\sigma$ .						

X(4350) BRANCHING RATIOS

$\Gamma(J/\psi\phi)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
seen	<sup>1</sup> SHEN 10	BELL	10.6 $e^+ e^- \rightarrow e^+ e^- J/\psi\phi$		
<sup>1</sup> Statistical significance of 3.2 $\sigma$ .					
$\Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	VALUE	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
seen	<sup>1</sup> SHEN 10	BELL	10.6 $e^+ e^- \rightarrow e^+ e^- J/\psi\phi$		
<sup>1</sup> Statistical significance of 3.2 $\sigma$ .					

X(4350) REFERENCES

SHEN 10	PRL 104 112004	C.P. Shen <i>et al.</i>	(BELLE Collab.)
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X(4360)

$I^G(J^{PC}) = ?^?(1^{--})$

Seen in radiative return from  $e^+ e^-$  collisions at  $\sqrt{s} = 9.54\text{--}10.58$  GeV by AUBERT 07S, WANG 07D, and LEES 14F. See also the review under the X(3872) particle listings. (See the index for the page number.)

X(4360) MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4346 ± 6 OUR AVERAGE</b>				
4347 ± 6 ± 3	279	<sup>1</sup> WANG 15A	BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
4340 ± 16 ± 9	37	<sup>2</sup> LEES 14F	BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4355 <sup>+9</sup> <sub>-10</sub> ± 9	74	<sup>3</sup> LIU 08H	RVUE	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
4324 ± 24		<sup>4</sup> AUBERT 07s	BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
4361 ± 9 ± 9	47	<sup>2</sup> WANG 07D	BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
<sup>1</sup> From a two-resonance fit. Supersedes WANG 07D. <sup>2</sup> From a two-resonance fit. <sup>3</sup> From a combined fit of AUBERT 07s and WANG 07D data with two resonances. <sup>4</sup> From a single-resonance fit. Systematic errors not estimated.				

X(4360) WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>102 ± 10 OUR AVERAGE</b>				
103 ± 9 ± 5	279	<sup>1</sup> WANG 15A	BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
94 ± 32 ± 13	37	<sup>2</sup> LEES 14F	BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
103 <sup>+17</sup> <sub>-15</sub> ± 11	74	<sup>3</sup> LIU 08H	RVUE	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
172 ± 33		<sup>4</sup> AUBERT 07s	BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
74 ± 15 ± 10	47	<sup>2</sup> WANG 07D	BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
<sup>1</sup> From a two-resonance fit. Supersedes WANG 07D. <sup>2</sup> From a two-resonance fit. <sup>3</sup> From a combined fit of AUBERT 07s and WANG 07D data with two resonances. <sup>4</sup> From a single-resonance fit. Systematic errors not estimated.				

X(4360) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $e^+ e^-$	
$\Gamma_2$ $\psi(2S) \pi^+ \pi^-$	seen
$\Gamma_3$ $\psi(3823) \pi^+ \pi^-$	possibly seen
$\Gamma_4$ $J/\psi \eta$	
$\Gamma_5$ $D^0 D^{*-} \pi^+$	
$\Gamma_6$ $\chi_{c1} \gamma$	
$\Gamma_7$ $\chi_{c2} \gamma$	

Meson Particle Listings

X(4360),  $\psi(4415)$

X(4360) $\Gamma(i) \times \Gamma(e^+ e^-)/\Gamma(\text{total})$					$\Gamma_2\Gamma_1/\Gamma$
$\Gamma(\psi(2S)\pi^+\pi^-) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$9.2 \pm 0.6 \pm 0.6$	279	<sup>1</sup> WANG	15A	BELL	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$10.9 \pm 0.6 \pm 0.7$	279	<sup>2</sup> WANG	15A	BELL	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$6.0 \pm 1.0 \pm 0.5$	37	<sup>3</sup> LEES	14F	BABR	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$7.2 \pm 1.0 \pm 0.6$	37	<sup>4</sup> LEES	14F	BABR	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$11.1^{+1.3}_{-1.2}$	74	<sup>5</sup> LIU	08H	RVUE	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$12.3 \pm 1.2$	74	<sup>6</sup> LIU	08H	RVUE	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$10.4 \pm 1.7 \pm 1.5$	47	<sup>3</sup> WANG	07D	BELL	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
$11.8 \pm 1.8 \pm 1.4$	47	<sup>4</sup> WANG	07D	BELL	$10.58 e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$

- <sup>1</sup> Solution I of two equivalent solutions from a fit using two interfering resonances. Supercedes WANG 07D.  
<sup>2</sup> Solution II of two equivalent solutions from a fit using two interfering resonances. Supercedes WANG 07D.  
<sup>3</sup> Solution I of two equivalent solutions in a fit using two interfering resonances.  
<sup>4</sup> Solution II of two equivalent solutions in a fit using two interfering resonances.  
<sup>5</sup> Solution I in a combined fit of AUBERT 07s and WANG 07D data with two resonances.  
<sup>6</sup> Solution II in a combined fit of AUBERT 07s and WANG 07D data with two resonances.

$\Gamma(J/\psi\eta) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_4\Gamma_1/\Gamma$
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<6.8	90	WANG	13B	BELL	$e^+e^- \rightarrow J/\psi\eta\gamma$

$\Gamma(\chi_{c1}\gamma) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_6\Gamma_1/\Gamma$	
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;0.57</b>	90	<sup>1</sup> HAN	15	BELL	$10.58 e^+e^- \rightarrow \chi_{c1}\gamma$	
<sup>1</sup> Using $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.21)\%$ .						

$\Gamma(\chi_{c2}\gamma) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_7\Gamma_1/\Gamma$	
VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;1.9</b>	90	<sup>1</sup> HAN	15	BELL	$10.58 e^+e^- \rightarrow \chi_{c2}\gamma$	
<sup>1</sup> Using $B(\eta \rightarrow \gamma\gamma) = (39.41 \pm 0.21)\%$ .						

X(4360) BRANCHING RATIOS

$\Gamma(D^0D^{*-}\pi^+)/\Gamma(\psi(2S)\pi^+\pi^-)$					$\Gamma_5/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;8</b>	90	PAKHLOVA	09	BELL	$e^+e^- \rightarrow X(4360) \rightarrow D^0D^{*-}\pi^+$	

$\Gamma(\psi(3823)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
possibly seen	19	<sup>1</sup> ABLIKIM	15s	BES3	$e^+e^- \rightarrow \pi^+\pi^-\chi_{c1}\gamma$	

- <sup>1</sup> From a fit of  $e^+e^- \rightarrow \pi^+\pi^-\psi(3823)$ ,  $\psi(3823) \rightarrow \chi_{c1}\gamma$  cross sections taken at  $\sqrt{s}$  values of 4.23, 4.26, 4.36, 4.42, and 4.60 GeV to the X(4360) line shape.

$\Gamma(D^0D^{*-}\pi^+)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma \times \Gamma_1/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;0.72 <math>\times 10^{-6}</math></b>	90	<sup>1</sup> PAKHLOVA	09	BELL	$e^+e^- \rightarrow X(4360) \rightarrow D^0D^{*-}\pi^+$	
<sup>1</sup> Using $4355^{+9}_{-10} \pm 9$ MeV for the mass of X(4360).						

X(4360) REFERENCES

ABLIKIM	15S	PRL 115 011803	M. Ablikim <i>et al.</i>	(BES III Collab.)
HAN	15	PR D92 012011	Y.-L. Han <i>et al.</i>	(BELLE Collab.)
WANG	15A	PR D91 112007	X.-L. Wang <i>et al.</i>	(BELLE Collab.)
LEES	14F	PR D89 111103	J.P. Lees <i>et al.</i>	(BABAR Collab.)
WANG	13B	PR D87 051101	X.-L. Wang <i>et al.</i>	(BELLE Collab.)
PAKHLOVA	09	PR D80 091101	G. Pakhlova <i>et al.</i>	(BELLE Collab.)
LIU	08H	PR D78 014032	Z.Q. Liu, X.S. Qin, C.Z. Yuan	
AUBERT	07S	PRL 98 212001	B. Aubert <i>et al.</i>	(BABAR Collab.)
WANG	07D	PRL 99 142002	X.-L. Wang <i>et al.</i>	(BELLE Collab.)

$$\psi(4415)$$

$$I^G(J^{PC}) = 0^-(1^--)$$

$\psi(4415)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4421 <math>\pm</math> 4 OUR ESTIMATE</b>			
<b>4415.1 <math>\pm</math> 7.9</b>	<sup>1</sup> ABLIKIM	08D	BES2 $e^+e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4412 $\pm$ 15	<sup>2</sup> MO	10	RVUE $e^+e^- \rightarrow$ hadrons
4411 $\pm$ 7	<sup>3</sup> PAKHLOVA	08A	BELL $10.6 e^+e^- \rightarrow D^0D^-\pi^+\gamma$
4425 $\pm$ 6	<sup>4</sup> SETH	05A	RVUE $e^+e^- \rightarrow$ hadrons
4429 $\pm$ 9	<sup>5</sup> SETH	05A	RVUE $e^+e^- \rightarrow$ hadrons
4417 $\pm$ 10	BRANDELIK	78C	DASP $e^+e^-$
4414 $\pm$ 7	SIEGRIST	76	MRK1 $e^+e^-$

- <sup>1</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (234 \pm 88)^\circ$ .  
<sup>2</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects.  
<sup>3</sup> Systematic uncertainties not estimated.  
<sup>4</sup> From a fit to Crystal Ball (OSTERHELD 86) data.  
<sup>5</sup> From a fit to BES (BAI 02c) data.

$\psi(4415)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>62 <math>\pm</math> 20 OUR ESTIMATE</b>			
<b>71.5 <math>\pm</math> 19.0</b>	<sup>6</sup> ABLIKIM	08D	BES2 $e^+e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
118 $\pm$ 32	<sup>7</sup> MO	10	RVUE $e^+e^- \rightarrow$ hadrons
77 $\pm$ 20	<sup>8</sup> PAKHLOVA	08A	BELL $10.6 e^+e^- \rightarrow D^0D^-\pi^+\gamma$
119 $\pm$ 16	<sup>9</sup> SETH	05A	RVUE $e^+e^- \rightarrow$ hadrons
118 $\pm$ 35	<sup>10</sup> SETH	05A	RVUE $e^+e^- \rightarrow$ hadrons
66 $\pm$ 15	BRANDELIK	78C	DASP $e^+e^-$
33 $\pm$ 10	SIEGRIST	76	MRK1 $e^+e^-$

- <sup>6</sup> Reanalysis of data presented in BAI 02c. From a global fit over the center-of-mass energy region 3.7–5.0 GeV covering the  $\psi(3770)$ ,  $\psi(4040)$ ,  $\psi(4160)$ , and  $\psi(4415)$  resonances. Phase angle fixed in the fit to  $\delta = (234 \pm 88)^\circ$ .  
<sup>7</sup> Reanalysis of data presented in BAI 00 and BAI 02c. From a global fit over the center-of-mass energy 3.8–4.8 GeV covering the  $\psi(4040)$ ,  $\psi(4160)$  and  $\psi(4415)$  resonances and including interference effects.  
<sup>8</sup> Systematic uncertainties not estimated.  
<sup>9</sup> From a fit to Crystal Ball (OSTERHELD 86) data.  
<sup>10</sup> From a fit to BES (BAI 02c) data.

$\psi(4415)$  DECAY MODES

Due to the complexity of the  $c\bar{c}$  threshold region, in this listing, “seen” (“not seen”) means that a cross section for the mode in question has been measured at effective  $\sqrt{s}$  near this particle’s central mass value, more (less) than  $2\sigma$  above zero, without regard to any peaking behavior in  $\sqrt{s}$  or absence thereof. See mode listing(s) for details and references.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $D\bar{D}$	seen	
$\Gamma_2$ $D^0\bar{D}^0$	seen	
$\Gamma_3$ $D^+D^-$	seen	
$\Gamma_4$ $D^*\bar{D} + \text{c.c.}$	seen	
$\Gamma_5$ $D^*(2007)^0\bar{D}^0 + \text{c.c.}$	seen	
$\Gamma_6$ $D^*(2010)^+D^- + \text{c.c.}$	seen	
$\Gamma_7$ $D^*\bar{D}^*$	seen	
$\Gamma_8$ $D^*(2007)^0\bar{D}^*(2007)^0 + \text{c.c.}$	seen	
$\Gamma_9$ $D^*(2010)^+D^*(2010)^- + \text{c.c.}$	seen	
$\Gamma_{10}$ $D^0D^-\pi^+$ (excl. $D^*(2007)^0\bar{D}^0 + \text{c.c., } D^*(2010)^+D^- + \text{c.c.}$ )	$< 2.3$ %	90%
$\Gamma_{11}$ $D\bar{D}_s^*(2460) \rightarrow D^0D^-\pi^+ + \text{c.c.}$	$(10 \pm 4)$ %	
$\Gamma_{12}$ $D^0\bar{D}^{*-}\pi^+ + \text{c.c.}$	$< 11$ %	90%
$\Gamma_{13}$ $D_s^+D_s^-$	not seen	
$\Gamma_{14}$ $\omega\chi_{c2}$	possibly seen	
$\Gamma_{15}$ $D_s^{*+}D_s^- + \text{c.c.}$	seen	
$\Gamma_{16}$ $D_s^{*+}D_s^{*-}$	not seen	
$\Gamma_{17}$ $\psi(3823)\pi^+\pi^-$	possibly seen	
$\Gamma_{18}$ $J/\psi\eta$	$< 6$ $\times 10^{-3}$	90%
$\Gamma_{19}$ $\chi_{c1}\gamma$	$< 8$ $\times 10^{-4}$	90%
$\Gamma_{20}$ $\chi_{c2}\gamma$	$< 4$ $\times 10^{-3}$	90%
$\Gamma_{21}$ $e^+e^-$	$(9.4 \pm 3.2) \times 10^{-6}$	





Meson Particle Listings

X(4430)<sup>±</sup>, X(4660)

X(4430) <sup>±</sup> WIDTH				
VALUE (MeV)	DOCUMENT ID		TECN	COMMENT
<b>181±31 OUR AVERAGE</b>				
172+13+37 -34	<sup>1</sup> AAIJ	14AG LHCb	$B^0 \rightarrow K^+ \pi^- \psi(2S)$	
200+41+26 -46-35	<sup>1</sup> CHILIKIN	13 BELL	$B^0 \rightarrow K^+ \pi^- \psi(2S)$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
107+86+74 -43-56	<sup>2</sup> MIZUK	09 BELL	$B \rightarrow K \pi^+ \psi(2S)$	
45+18+30 -13-13	<sup>3</sup> CHOI	08 BELL	$B \rightarrow K \pi^+ \psi(2S)$	
<sup>1</sup> From a four-dimensional amplitude analysis. <sup>2</sup> From a Dalitz plot analysis. Superseded by CHILIKIN 13. <sup>3</sup> Superseded by MIZUK 09 and CHILIKIN 13.				

X(4430) <sup>±</sup> DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \quad \pi^+ \psi(2S)$	seen	
$\Gamma_2 \quad \pi^+ J/\psi$	seen	

X(4430) <sup>±</sup> BRANCHING RATIOS				
Γ(π <sup>+</sup> ψ(2S))/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT	
seen	<sup>1</sup> AAIJ	14AG LHCb	B <sup>0</sup> → K <sup>+</sup> π <sup>-</sup> ψ(2S)	
seen	<sup>2</sup> CHILIKIN	13 BELL	B <sup>0</sup> → K <sup>+</sup> π <sup>-</sup> ψ(2S)	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
not seen	<sup>3</sup> AUBERT	09AA BABR	B → K π <sup>+</sup> ψ(2S)	
seen	<sup>4</sup> MIZUK	09 BELL	B → K π <sup>+</sup> ψ(2S)	
<sup>1</sup> From a four-dimensional amplitude analysis. No product of branching fractions quoted.				
<sup>2</sup> From a four-dimensional amplitude analysis. Measured a product of branching fractions B(B <sup>0</sup> → X(4430) <sup>-</sup> K <sup>+</sup> )×B(X(4430) <sup>-</sup> → ψ(2S) π <sup>-</sup> ) = (6.0 <sup>+1.7+2.5</sup> <sub>-2.0-1.4</sub> ) × 10 <sup>-5</sup> .				
<sup>3</sup> AUBERT 09AA quotes B(B <sup>+</sup> → K <sup>0</sup> X(4430) <sup>+</sup> ) × B(X(4430) <sup>+</sup> → π <sup>+</sup> ψ(2S)) < 4.7×10 <sup>-5</sup> and B(B <sup>0</sup> → K <sup>-</sup> X(4430) <sup>+</sup> ) × B(X(4430) <sup>+</sup> → π <sup>+</sup> ψ(2S)) < 3.1×10 <sup>-5</sup> at 95% CL.				
<sup>4</sup> Measured a product of branching fractions B(B <sup>0</sup> → K <sup>-</sup> X(4430) <sup>+</sup> ) × B(X(4430) <sup>+</sup> → π <sup>+</sup> ψ(2S)) = (3.2 <sup>+1.8+5.3</sup> <sub>-0.9-1.6</sub> ) × 10 <sup>-5</sup> . Superseded by CHILIKIN 13.				

$\Gamma(\pi^+ J/\psi)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	<sup>1</sup> CHILIKIN	14 BELL	$\bar{B}^0 \rightarrow K^- \pi^+ J/\psi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	<sup>2</sup> AUBERT	09AA BABR	$B \rightarrow K \pi^+ J/\psi$
<sup>1</sup> CHILIKIN 14 reports $B(\bar{B}^0 \rightarrow X(4430)^+ K^-) \times B(X(4430)^+ \rightarrow J/\psi \pi^+) = (5.4^{+4.0+1.1}_{-1.0-0.9}) \times 10^{-6}$ .			
<sup>2</sup> AUBERT 09AA quotes $B(B^+ \rightarrow \bar{K}^0 X(4430)^+) \times B(X(4430)^+ \rightarrow \pi^+ J/\psi) < 1.5 \times 10^{-5}$ and $B(\bar{B}^0 \rightarrow K^- X(4430)^+) \times B(X(4430)^+ \rightarrow \pi^+ J/\psi) < 0.4 \times 10^{-5}$ at 95% CL.			

X(4430) <sup>±</sup> REFERENCES				
AAIJ	15BH	PR D92 112009	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14AG	PRL 112 222002	R. Aaij <i>et al.</i>	(LHCb Collab.) JP
CHILIKIN	14	PR D90 112009	K. Chilikin <i>et al.</i>	(BELLE Collab.)
CHILIKIN	13	PR D88 074026	K. Chilikin <i>et al.</i>	(BELLE Collab.) JP
AUBERT	09AA	PR D79 112001	B. Aubert <i>et al.</i>	(BABAR Collab.)
MIZUK	09	PR D80 031104	R. Mizuk <i>et al.</i>	(BELLE Collab.)
CHOI	08	PRL 100 142001	S.-K. Choi <i>et al.</i>	(BELLE Collab.)

X(4660)	$i^G(J^{PC}) = ?^?(1^{--})$
Seen in radiative return from $e^+ e^-$ collisions at $\sqrt{s} = 9.54\text{--}10.58$ GeV by WANG 07D. Also obtained in a combined fit of WANG 07D, AUBERT 07S, and LEES 14F. See also the review under the X(3872) particle listings. (See the index for the page number.)	

X(4660) MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4643± 9 OUR AVERAGE</b> Error includes scale factor of 1.2.				
4652±10±11	279	<sup>1</sup> WANG	15A BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
4669±21± 3	37	<sup>2</sup> LEES	14F BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
4634 <sup>+ 8</sup> <sub>- 7</sub>	142	<sup>3</sup> PAKHLOVA	08B BELL	$e^+ e^- \rightarrow \Lambda_c^+ \Lambda_c^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4661 <sup>+ 9</sup> <sub>- 8</sub>	6	<sup>4</sup> LIU	08H RVUE	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
4664±11± 5	44	WANG	07D BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$

- <sup>1</sup> From a two-resonance fit. Supersedes WANG 07D.  
<sup>2</sup> From a two-resonance fit.  
<sup>3</sup> The  $\pi^+ \pi^- \psi(2S)$  and  $\Lambda_c^+ \Lambda_c^-$  states are not necessarily the same.  
<sup>4</sup> From a combined fit of AUBERT 07S and WANG 07D data with two resonances.

X(4660) WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>72±11 OUR AVERAGE</b>				
68±11± 5	279	<sup>1</sup> WANG	15A BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
104±48±10	37	<sup>2</sup> LEES	14F BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
92±40 <sup>+10</sup> <sub>-24</sub>	142	<sup>3</sup> PAKHLOVA	08B BELL	$e^+ e^- \rightarrow \Lambda_c^+ \Lambda_c^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
42±17 <sup>+17</sup> <sub>-12</sub>	6	<sup>4</sup> LIU	08H RVUE	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
48±15± 3	44	WANG	07D BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$

- <sup>1</sup> From a two-resonance fit. Supersedes WANG 07D.  
<sup>2</sup> From a two-resonance fit.  
<sup>3</sup> The  $\pi^+ \pi^- \psi(2S)$  and  $\Lambda_c^+ \Lambda_c^-$  states are not necessarily the same.  
<sup>4</sup> From a combined fit of AUBERT 07S and WANG 07D data with two resonances.

X(4660) DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \quad e^+ e^-$		
$\Gamma_2 \quad \psi(2S) \pi^+ \pi^-$	seen	
$\Gamma_3 \quad J/\psi \eta$		
$\Gamma_4 \quad D^0 D^{*-} \pi^+$		
$\Gamma_5 \quad \chi_{c1} \gamma$		
$\Gamma_6 \quad \chi_{c2} \gamma$		
$\Gamma_7 \quad \Lambda_c^+ \Lambda_c^-$		

X(4660) $\Gamma(i) \times \Gamma(e^+ e^-)/\Gamma(\text{total})$				
$\Gamma(\psi(2S) \pi^+ \pi^-) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
2.0±0.3±0.2	279	<sup>1</sup> WANG	15A BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
8.1±1.1±1.0	279	<sup>2</sup> WANG	15A BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
2.7±1.3±0.5	37	<sup>3</sup> LEES	14F BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
7.5±1.7±0.7	37	<sup>4</sup> LEES	14F BABR	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
2.2 <sup>+0.7</sup> <sub>-0.6</sub>	44	<sup>5</sup> LIU	08H RVUE	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
5.9±1.6	44	<sup>6</sup> LIU	08H RVUE	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
3.0±0.9±0.3	44	<sup>3</sup> WANG	07D BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$
7.6±1.8±0.8	44	<sup>4</sup> WANG	07D BELL	10.58 $e^+ e^- \rightarrow \gamma \pi^+ \pi^- \psi(2S)$

- <sup>1</sup> Solution I of two equivalent solutions from a fit using two interfering resonances. Supersedes WANG 07D.  
<sup>2</sup> Solution II of two equivalent solutions from a fit using two interfering resonances. Supersedes WANG 07D.  
<sup>3</sup> Solution I of two equivalent solutions in a fit using two interfering resonances.  
<sup>4</sup> Solution II of two equivalent solutions in a fit using two interfering resonances.  
<sup>5</sup> Solution I in a combined fit of AUBERT 07S and WANG 07D data with two resonances.  
<sup>6</sup> Solution II in a combined fit of AUBERT 07S and WANG 07D data with two resonances.

$\Gamma(J/\psi \eta) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
VALUE (eV)	CL%			
<0.94	90	WANG	13B BELL	$e^+ e^- \rightarrow J/\psi \eta \gamma$

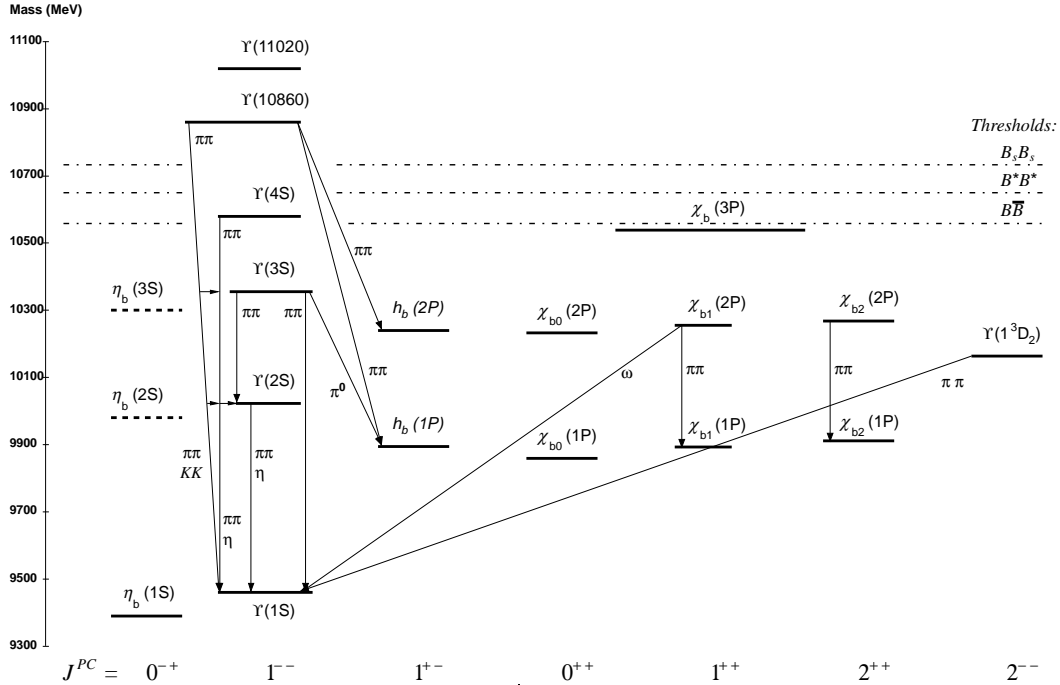
$\Gamma(\chi_{c1} \gamma) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
VALUE (eV)	CL%			
<0.45	90	<sup>1</sup> HAN	15 BELL	10.58 $e^+ e^- \rightarrow \chi_{c1} \gamma$
<sup>1</sup> Using $B(\eta \rightarrow \gamma \gamma) = (39.41 \pm 0.21)\%$ .				

$\Gamma(\chi_{c2} \gamma) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
VALUE (eV)	CL%			
<2.1	90	<sup>1</sup> HAN	15 BELL	10.58 $e^+ e^- \rightarrow \chi_{c2} \gamma$
<sup>1</sup> Using $B(\eta \rightarrow \gamma \gamma) = (39.41 \pm 0.21)\%$ .				

See key on page 601

Meson Particle Listings  
X(4660)

X(4660) BRANCHING RATIOS					X(4660) REFERENCES			
$\Gamma(D^0 D^{*-} \pi^+)/\Gamma(\psi(2S) \pi^+ \pi^-)$					$\Gamma_4/\Gamma_2$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT				
<10	90	PAKHLOVA	09	BELL	$e^+ e^- \rightarrow D^0 D^{*-} \pi^+$			
$\Gamma(D^0 D^{*-} \pi^+)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma \times \Gamma_1/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT				
<0.37 × 10 <sup>-6</sup>	90	<sup>1</sup> PAKHLOVA	09	BELL	$e^+ e^- \rightarrow D^0 D^{*-} \pi^+$			
<sup>1</sup> Using 4664 ± 11 ± 5 MeV for the mass of X(4660).								
$\Gamma(\Lambda_c^+ \Lambda_c^-)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma \times \Gamma_1/\Gamma$			
VALUE (units 10 <sup>-6</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT				
0.68 <sup>+0.16+0.29</sup> <sub>-0.15-0.30</sub>	142	<sup>1</sup> PAKHLOVA	08B	BELL	$e^+ e^- \rightarrow \Lambda_c^+ \Lambda_c^-$			
<sup>1</sup> The $\pi^+ \pi^- \psi(2S)$ and $\Lambda_c^+ \Lambda_c^-$ states are not necessarily the same.								
					HAN	15	PR D92 012011	Y.L. Han <i>et al.</i> (BELLE Collab.)
					WANG	15A	PR D91 112007	X.L. Wang <i>et al.</i> (BELLE Collab.)
					LEES	14F	PR D89 111103	J.P. Lees <i>et al.</i> (BABAR Collab.)
					WANG	13B	PR D87 051101	X.L. Wang <i>et al.</i> (BELLE Collab.)
					PAKHLOVA	09	PR D80 091101	G. Pakhlova <i>et al.</i> (BELLE Collab.)
					LIU	08H	PR D78 014032	Z.Q. Liu, X.S. Qin, C.Z. Yuan
					PAKHLOVA	08B	PRL 101 172001	C. Pakhlova <i>et al.</i> (BELLE Collab.)
					AUBERT	07S	PRL 98 212001	B. Aubert <i>et al.</i> (BABAR Collab.)
					WANG	07D	PRL 99 142002	X.L. Wang <i>et al.</i> (BELLE Collab.)

**$b\bar{b}$  MESONS**
**THE BOTTOMONIUM SYSTEM**


The level scheme of the  $b\bar{b}$  states showing experimentally established states with solid lines. Singlet states are called  $\eta_b$  and  $h_b$ , triplet states  $\Upsilon$  and  $\chi_{bJ}$ . In parentheses it is sufficient to give the radial quantum number and the orbital angular momentum to specify the states with all their quantum numbers. *E.g.*,  $h_b(2P)$  means  $2^1P_1$  with  $n = 2$ ,  $L = 1$ ,  $S = 0$ ,  $J = 1$ ,  $PC = +-$ . The figure shows observed hadronic transitions. The single photon transitions  $\Upsilon(nS) \rightarrow \gamma\eta_b(mS)$ ,  $\Upsilon(nS) \rightarrow \gamma\chi_{bJ}(mP)$ , and  $\chi_{bJ}(nP) \rightarrow \gamma\Upsilon(mS)$  are omitted for clarity.

**WIDTH DETERMINATIONS OF THE  $\Upsilon$  STATES**

As is the case for the  $J/\psi(1S)$  and  $\psi(2S)$ , the full widths of the  $b\bar{b}$  states  $\Upsilon(1S)$ ,  $\Upsilon(2S)$ , and  $\Upsilon(3S)$  are not directly measurable, since they are much narrower than the energy resolution of the  $e^+e^-$  storage rings where these states are produced. The common indirect method to determine  $\Gamma$  starts from

$$\Gamma = \Gamma_{\ell\ell}/B_{\ell\ell} \ , \quad (1)$$

where  $\Gamma_{\ell\ell}$  is one leptonic partial width and  $B_{\ell\ell}$  is the corresponding branching fraction ( $\ell = e, \mu$ , or  $\tau$ ). One then assumes  $e$ - $\mu$ - $\tau$  universality and uses

$$\begin{aligned} \Gamma_{\ell\ell} &= \Gamma_{ee} \\ B_{\ell\ell} &= \text{average of } B_{ee}, B_{\mu\mu}, \text{ and } B_{\tau\tau} \ . \end{aligned} \quad (2)$$

The electronic partial width  $\Gamma_{ee}$  is also not directly measurable at  $e^+e^-$  storage rings, only in the combination  $\Gamma_{ee}\Gamma_{\text{had}}/\Gamma$ , where  $\Gamma_{\text{had}}$  is the hadronic partial width and

$$\Gamma_{\text{had}} + 3\Gamma_{ee} = \Gamma \ . \quad (3)$$

This combination is obtained experimentally from the energy-integrated hadronic cross section

$$\begin{aligned} \int_{\text{resonance}} \sigma(e^+e^- \rightarrow \Upsilon \rightarrow \text{hadrons})dE \\ = \frac{6\pi^2 \Gamma_{ee}\Gamma_{\text{had}}}{M^2 \Gamma} C_r = \frac{6\pi^2 \Gamma_{ee}^{(0)}\Gamma_{\text{had}}}{M^2 \Gamma} C_r^{(0)} \ , \end{aligned} \quad (4)$$

where  $M$  is the  $\Upsilon$  mass, and  $C_r$  and  $C_r^{(0)}$  are radiative correction factors.  $C_r$  is used for obtaining  $\Gamma_{ee}$  as defined in Eq. (1), and contains corrections from all orders of QED for describing  $(b\bar{b}) \rightarrow e^+e^-$ . The lowest order QED value  $\Gamma_{ee}^{(0)}$ , relevant for comparison with potential-model calculations, is defined by the lowest order QED graph (Born term) alone, and is about 7% lower than  $\Gamma_{ee}$ .

The Listings give experimental results on  $B_{ee}$ ,  $B_{\mu\mu}$ ,  $B_{\tau\tau}$ , and  $\Gamma_{ee}\Gamma_{\text{had}}/\Gamma$ . The entries of the last quantity have been re-evaluated consistently using the correction procedure of KU-RAEV 85. The partial width  $\Gamma_{ee}$  is obtained from the average values for  $\Gamma_{ee}\Gamma_{\text{had}}/\Gamma$  and  $B_{\ell\ell}$  using

$$\Gamma_{ee} = \frac{\Gamma_{ee}\Gamma_{\text{had}}}{\Gamma(1 - 3B_{\ell\ell})} \ . \quad (5)$$

See key on page 601

# Meson Particle Listings

## Bottomonium, $\eta_b(1S)$

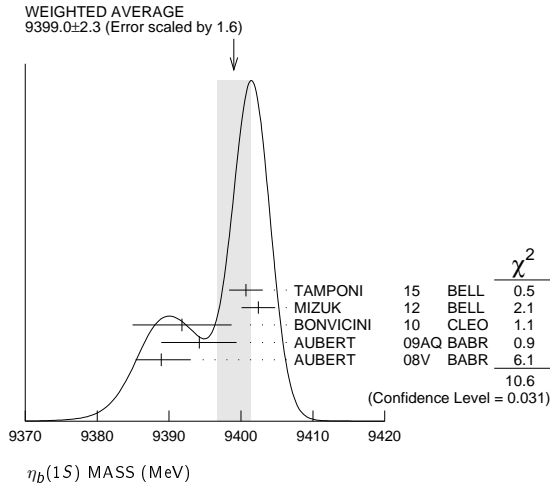
The total width  $\Gamma$  is then obtained from Eq. (1). We do not list  $\Gamma_{ee}$  and  $\Gamma$  values of individual experiments. The  $\Gamma_{ee}$  values in the Meson Summary Table are also those defined in Eq. (1).

<b><math>\eta_b(1S)</math></b>	$I^G(J^{PC}) = 0^+(0^{-+})$
Quantum numbers shown are quark-model predictions. Observed in radiative decay of the $\Upsilon(3S)$ , therefore $C = +$ .	

### $\eta_b(1S)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9399.0 ± 2.3 OUR AVERAGE</b>		Error includes scale factor of 1.6. See the ideogram below.		
9400.7 ± 1.7 ± 1.6	33.1k	TAMPONI	15	BELL $e^+e^- \rightarrow \gamma\eta + \text{hadrons}$
9402.4 ± 1.5 ± 1.8	34k	<sup>1</sup> MIZUK	12	BELL $e^+e^- \rightarrow \gamma\pi^+\pi^- + \text{hadrons}$
9391.8 ± 6.6 ± 2.0	2.3k	<sup>2</sup> BONVICINI	10	CLEO $\Upsilon(3S) \rightarrow \gamma X$
9394.2 ± 4.8 ± 2.0	13k	<sup>2</sup> AUBERT	09AQ	BABR $\Upsilon(2S) \rightarrow \gamma X$
9388.9 ± 3.1 ± 2.7	19k	<sup>2</sup> AUBERT	08V	BABR $\Upsilon(3S) \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
9393.2 ± 3.4 ± 2.3	10	<sup>2,3</sup> DOBBS	12	$\Upsilon(2S) \rightarrow \gamma \text{hadrons}$
9300 ± 20 ± 20		HEISTER	02D	ALEP 181-209 $e^+e^-$

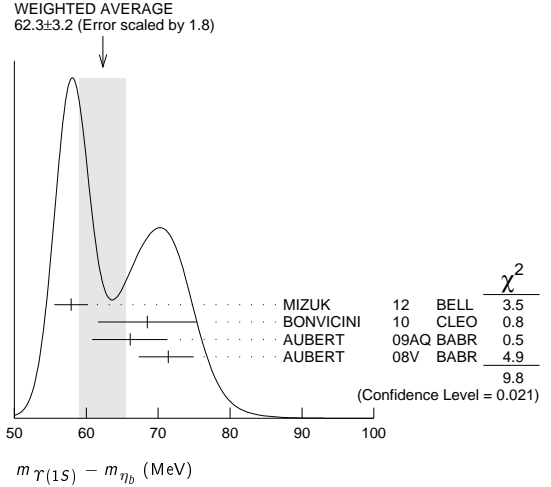
<sup>1</sup> With floating width. Not independent of the corresponding mass difference measurement.  
<sup>2</sup> Assuming  $\Gamma_{\eta_b(1S)} = 10$  MeV. Not independent of the corresponding  $\gamma$  energy or mass difference measurements.  
<sup>3</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.



### $m\Upsilon(1S) - m_{\eta_b}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>62.3 ± 3.2 OUR AVERAGE</b>		Error includes scale factor of 1.8. See the ideogram below.		
57.9 ± 1.5 ± 1.8	34k	<sup>4</sup> MIZUK	12	BELL $e^+e^- \rightarrow \gamma\pi^+\pi^- + \text{hadrons}$
68.5 ± 6.6 ± 2.0	2.3 ± 0.5k	<sup>5</sup> BONVICINI	10	CLEO $\Upsilon(3S) \rightarrow \gamma X$
66.1 ± 4.8 ± 2.0	13 ± 5k	<sup>5</sup> AUBERT	09AQ	BABR $\Upsilon(2S) \rightarrow \gamma X$
71.4 ± 2.3 ± 2.7	19 ± 3k	<sup>5</sup> AUBERT	08V	BABR $\Upsilon(3S) \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
67.1 ± 3.4 ± 2.3	10 ± 5	<sup>5,6</sup> DOBBS	12	$\Upsilon(2S) \rightarrow \gamma \text{hadrons}$

<sup>4</sup> With floating width. Not independent of the corresponding mass measurement.  
<sup>5</sup> Assuming  $\Gamma_{\eta_b(1S)} = 10$  MeV. Not independent of the corresponding  $\gamma$  energy or mass measurements.  
<sup>6</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.



### $\gamma$ ENERGY IN $\Upsilon(3S)$ DECAY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>920.6 ± 2.8 OUR AVERAGE</b>				
918.6 ± 6.0 ± 1.9	2.3 ± 0.5k	<sup>7</sup> BONVICINI	10	CLEO $\Upsilon(3S) \rightarrow \gamma X$
921.2 ± 2.1 ± 2.4	19 ± 3k	<sup>7</sup> AUBERT	08V	BABR $\Upsilon(3S) \rightarrow \gamma X$

<sup>7</sup> Assuming  $\Gamma_{\eta_b(1S)} = 10$  MeV. Not independent of the corresponding mass or mass difference measurements.

### $\gamma$ ENERGY IN $\Upsilon(2S)$ DECAY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>609.3 ± 4.6 ± 1.9</b>	13 ± 5k	<sup>8</sup> AUBERT	09AQ	BABR $\Upsilon(2S) \rightarrow \gamma X$

<sup>8</sup> Assuming  $\Gamma_{\eta_b(1S)} = 10$  MeV. Not independent of the corresponding mass or mass difference measurements.

### $\eta_b(1S)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10 ± 4 OUR AVERAGE</b>				
8 ± 6 ± 5	33.1k	<sup>9</sup> TAMPONI	15	BELL $e^+e^- \rightarrow \gamma\eta + \text{hadrons}$
10.8 ± 4.0 ± 4.5	34k	<sup>9</sup> MIZUK	12	BELL $e^+e^- \rightarrow \gamma\pi^+\pi^- + \text{hadrons}$

<sup>9</sup> With floating mass.

### $\eta_b(1S)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ hadrons	seen	
$\Gamma_2$ $3h^+3h^-$	not seen	
$\Gamma_3$ $2h^+2h^-$	not seen	
$\Gamma_4$ $4h^+4h^-$		
$\Gamma_5$ $\gamma\gamma$	not seen	
$\Gamma_6$ $\mu^+\mu^-$	$< 9 \times 10^{-3}$	90%
$\Gamma_7$ $\tau^+\tau^-$	$< 8\%$	90%

### $\eta_b(1S)$ $\Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
------------	-----	-------------	------	---------

• • • We do not use the following data for averages, fits, limits, etc. • • •

<470	95	ABDALLAH	06	DLPH 161-209 $e^+e^-$
<132	95	HEISTER	02D	ALEP 181-209 $e^+e^-$

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
------------	-----	-------------	------	---------

• • • We do not use the following data for averages, fits, limits, etc. • • •

<190	95	ABDALLAH	06	DLPH 161-209 $e^+e^-$
< 48	95	HEISTER	02D	ALEP 181-209 $e^+e^-$









# Meson Particle Listings

## $\Upsilon(1S)$

$\Gamma(\text{Sum of 100 exclusive modes})/\Gamma_{\text{total}}$				$\Gamma_{48}/\Gamma$	
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>1.200 \pm 0.017</math></b>	19,20	DOBBS 12A		$\Upsilon(1S) \rightarrow \text{hadrons}$	
<sup>19</sup> DOBBS 12A presents individual exclusive branching fractions or upper limits for 100 modes of four to ten pions, kaons, or protons.					
<sup>20</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.					
$\Gamma(ggg, \gamma g g \rightarrow \bar{d} \text{ anything})/\Gamma(ggg, \gamma g g \rightarrow \text{anything})$					
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.36 \pm 0.23 \pm 0.25</math></b>	455	ASNER 07	CLEO	$e^+ e^- \rightarrow \bar{d} X$	
$\Gamma(\gamma \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{49}/\Gamma$
VALUE (units $10^{-5}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>6.3 \pm 1.2 \pm 1.3</math></b>		<sup>21</sup> ANASTASSOV 99	CLE2	$e^+ e^- \rightarrow \text{hadrons}$	
<sup>21</sup> For $m_{\pi\pi} > 1 \text{ GeV}$ .					
$\Gamma(\gamma \pi^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$
VALUE (units $10^{-5}$ )		DOCUMENT ID	TECN	COMMENT	
<b><math>1.7 \pm 0.6 \pm 0.3</math></b>		<sup>22</sup> ANASTASSOV 99	CLE2	$e^+ e^- \rightarrow \text{hadrons}$	
<sup>22</sup> For $m_{\pi\pi} > 1 \text{ GeV}$ .					
$\Gamma(\gamma \pi^0 \eta)/\Gamma_{\text{total}}$					$\Gamma_{51}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 2.4</math></b>	90	<sup>23</sup> BESSON 07A	CLEO	$e^+ e^- \rightarrow \Upsilon(1S)$	
<sup>23</sup> BESSON 07A obtained this limit for $0.7 < m_{\pi^0 \eta} < 3 \text{ GeV}$ .					
$\Gamma(\gamma K^+ K^-)/\Gamma_{\text{total}}$ ( $2 < m_{K^+ K^-} < 3 \text{ GeV}$ )					$\Gamma_{52}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>1.14 \pm 0.08 \pm 0.10</math></b>	90	ATHAR 06	CLE3	$\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
$\Gamma(\gamma p \bar{p})/\Gamma_{\text{total}}$ ( $2 < m_{p \bar{p}} < 3 \text{ GeV}$ )					$\Gamma_{53}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt; 0.6</math></b>	90	ATHAR 06	CLE3	$\Upsilon(1S) \rightarrow \gamma p \bar{p}$	
$\Gamma(\gamma 2h^+ 2h^-)/\Gamma_{\text{total}}$					$\Gamma_{54}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>7.0 \pm 1.1 \pm 1.0</math></b>	$80 \pm 12$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 3h^+ 3h^-)/\Gamma_{\text{total}}$					$\Gamma_{55}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>5.4 \pm 1.5 \pm 1.3</math></b>	$39 \pm 11$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 4h^+ 4h^-)/\Gamma_{\text{total}}$					$\Gamma_{56}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>7.4 \pm 2.5 \pm 2.5</math></b>	$36 \pm 12$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma \pi^+ \pi^- K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{57}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.9 \pm 0.7 \pm 0.6</math></b>	$29 \pm 8$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 2\pi^+ 2\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{58}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.5 \pm 0.7 \pm 0.5</math></b>	$26 \pm 7$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 3\pi^+ 3\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{59}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.5 \pm 0.9 \pm 0.8</math></b>	$17 \pm 5$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{60}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.4 \pm 0.9 \pm 0.8</math></b>	$18 \pm 7$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma \pi^+ \pi^- p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{61}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.5 \pm 0.5 \pm 0.3</math></b>	$22 \pm 6$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 2\pi^+ 2\pi^- p \bar{p})/\Gamma_{\text{total}}$					$\Gamma_{62}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.4 \pm 0.4 \pm 0.4</math></b>	$7 \pm 6$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	
$\Gamma(\gamma 2K^+ 2K^-)/\Gamma_{\text{total}}$					$\Gamma_{63}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.2 \pm 0.2</math></b>	$2 \pm 2$	FULTON 90b	CLEO	$e^+ e^- \rightarrow \text{hadrons}$	

$\Gamma(\gamma\eta'(958))/\Gamma_{\text{total}}$					$\Gamma_{64}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 1.9	90	ATHAR 07A	CLEO	$\Upsilon(1S) \rightarrow \gamma\eta' \rightarrow \gamma\pi^+\pi^-\eta, \gamma\rho$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<16	90	RICHICHI 01B	CLE2	$\Upsilon(1S) \rightarrow \gamma\eta' \rightarrow \gamma\eta\pi^+\pi^-$	
$\Gamma(\gamma\eta)/\Gamma_{\text{total}}$					$\Gamma_{65}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 1.0	90	ATHAR 07A	CLEO	$\Upsilon(1S) \rightarrow \gamma\eta \rightarrow \gamma\gamma\gamma, \gamma\pi^+\pi^-\pi^0, \gamma 3\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<21	90	MASEK 02	CLEO	$\Upsilon(1S) \rightarrow \gamma\eta$	
$\Gamma(\gamma f_0(980))/\Gamma_{\text{total}}$					$\Gamma_{66}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<3	90	<sup>24</sup> ATHAR 06	CLE3	$\Upsilon(1S) \rightarrow \gamma\pi^+\pi^-$	
<sup>24</sup> Assuming $B(f_0(980) \rightarrow \pi\pi) = 1$ .					
$\Gamma(\gamma f_2'(1525))/\Gamma_{\text{total}}$					$\Gamma_{67}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.8<math>\pm</math>0.9 OUR AVERAGE</b>					
4.0 $\pm$ 1.4 $\pm$ 0.1	17 $\pm$ 5	<sup>25</sup> BESSON 11	CLEO	$\Upsilon(1S) \rightarrow K_S^0 K_S^0$	
3.7 $\pm$ <sup>+0.9</sup> <sub>-0.7</sub> $\pm$ 0.8		ATHAR 06	CLE3	$\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<14	90	<sup>26</sup> FULTON 90B	CLEO	$\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
<19.4	90	<sup>26</sup> ALBRECHT 89	ARG	$\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
<sup>25</sup> BESSON 11 reports $(4.0 \pm 1.3 \pm 0.6) \times 10^{-5}$ from a measurement of $[\Gamma(\Upsilon(1S) \rightarrow \gamma f_2'(1525))/\Gamma_{\text{total}}] \times [B(f_2'(1525) \rightarrow K\bar{K})]$ assuming $B(f_2'(1525) \rightarrow K\bar{K}) = (88.8 \pm 3.1) \times 10^{-2}$ , which we rescale to our best value $B(f_2'(1525) \rightarrow K\bar{K}) = (88.7 \pm 2.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. The result also assumes $B(K_S^0 \rightarrow \pi^+\pi^-) = (69.20 \pm 0.05)\%$ and $B(f_2'(1525) \rightarrow K\bar{K}) = 4 B(f_2'(1525) \rightarrow K_S^0 K_S^0)$ .					
<sup>26</sup> Assuming $B(f_2'(1525) \rightarrow K\bar{K}) = 0.71$ .					
$\Gamma(\gamma f_2(1270))/\Gamma_{\text{total}}$					$\Gamma_{68}/\Gamma$
VALUE (units $10^{-5}$ )	CL%		DOCUMENT ID	TECN	COMMENT
<b>10.1<math>\pm</math>0.9 OUR AVERAGE</b>					
10.5 $\pm$ 1.6 $\pm$ <sup>+1.9</sup> <sub>-1.8</sub>		<sup>27</sup> BESSON 07A	CLE3	$\Upsilon(1S) \rightarrow \gamma\pi^0\pi^0$	
10.2 $\pm$ 0.8 $\pm$ 0.7		ATHAR 06	CLE3	$\Upsilon(1S) \rightarrow \gamma\pi^+\pi^-$	
8.1 $\pm$ 2.3 $\pm$ <sup>+2.9</sup> <sub>-2.7</sub>		<sup>28</sup> ANASTASSOV 99	CLE2	$e^+e^- \rightarrow \text{hadrons}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<21	90	<sup>28</sup> FULTON 90B	CLEO	$\Upsilon(1S) \rightarrow \gamma\pi^+\pi^-$	
<13	90	<sup>28</sup> ALBRECHT 89	ARG	$\Upsilon(1S) \rightarrow \gamma\pi^+\pi^-$	
<81	90	SCHMITT 88	CBAL	$\Upsilon(1S) \rightarrow \gamma X$	
<sup>27</sup> Using $B(f_2(1270) \rightarrow \pi^0\pi^0) = B(f_2(1270) \rightarrow \pi\pi)/3$ and $B(f_2(1270) \rightarrow \pi\pi) = (0.845 \pm 0.025)\%$ .					
<sup>28</sup> Using $B(f_2(1270) \rightarrow \pi\pi) = 0.84$ .					
$\Gamma(\gamma\eta(1405))/\Gamma_{\text{total}}$					$\Gamma_{69}/\Gamma$
VALUE (units $10^{-5}$ )	CL%		DOCUMENT ID	TECN	COMMENT
<8.2	90	<sup>29</sup> FULTON 90B	CLEO	$\Upsilon(1S) \rightarrow \gamma K^\pm\pi^\mp K_S^0$	
<sup>29</sup> Includes unknown branching ratio of $\eta(1405) \rightarrow K^\pm\pi^\mp K_S^0$ .					
$\Gamma(\gamma f_0(1500))/\Gamma_{\text{total}}$					$\Gamma_{70}/\Gamma$
VALUE (units $10^{-5}$ )	CL%		DOCUMENT ID	TECN	COMMENT
<1.5	90	<sup>30</sup> BESSON 07A	CLEO	$e^+e^- \rightarrow \Upsilon(1S) \rightarrow \gamma\pi^0\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<6.1	90	<sup>31</sup> BESSON 07A	CLEO	$e^+e^- \rightarrow \Upsilon(1S) \rightarrow \gamma\eta\eta$	
<sup>30</sup> Using $B(f_0(1500) \rightarrow \pi^0\pi^0) = B(f_0(1500) \rightarrow \pi\pi)/3$ and $B(f_0(1500) \rightarrow \pi\pi) = (0.349 \pm 0.023)\%$ .					
<sup>31</sup> Calculated by us using $B(f_0(1500) \rightarrow \eta\eta) = (5.1 \pm 0.9)\%$ .					
$\Gamma(\gamma f_0(1710))/\Gamma_{\text{total}}$					$\Gamma_{71}/\Gamma$
VALUE (units $10^{-4}$ )	CL%		DOCUMENT ID	TECN	COMMENT
< 2.6	90	<sup>32</sup> ALBRECHT 89	ARG	$\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 6.3	90	<sup>32</sup> FULTON 90B	CLEO	$\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
<19	90	<sup>32</sup> FULTON 90B	CLEO	$\Upsilon(1S) \rightarrow \gamma K_S^0 K_S^0$	
< 8	90	<sup>33</sup> ALBRECHT 89	ARG	$\Upsilon(1S) \rightarrow \gamma\pi^+\pi^-$	
<24	90	<sup>34</sup> SCHMITT 88	CBAL	$\Upsilon(1S) \rightarrow \gamma X$	
<sup>32</sup> Assuming $B(f_0(1710) \rightarrow K\bar{K}) = 0.38$ .					
<sup>33</sup> Assuming $B(f_0(1710) \rightarrow \pi\pi) = 0.04$ .					
<sup>34</sup> Assuming $B(f_0(1710) \rightarrow \eta\eta) = 0.18$ .					

# Meson Particle Listings

## $\Upsilon(1S)$

$\Gamma(\gamma f_0(1710) \rightarrow \gamma K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{72}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<0.7	90	ATHAR	06	CLEO $e^+e^- \rightarrow \Upsilon(1S) \rightarrow \gamma K^+ K^-$	

$\Gamma(\gamma f_0(1710) \rightarrow \gamma \pi^0 \pi^0)/\Gamma_{\text{total}}$					$\Gamma_{73}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.4	90	BESSON	07A	CLEO $e^+e^- \rightarrow \Upsilon(1S) \rightarrow \gamma \pi^0 \pi^0$	

$\Gamma(\gamma f_0(1710) \rightarrow \gamma \eta \eta)/\Gamma_{\text{total}}$					$\Gamma_{74}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.8	90	BESSON	07A	CLEO $e^+e^- \rightarrow \Upsilon(1S) \rightarrow \gamma \eta \eta$	

$\Gamma(\gamma f_4(2050))/\Gamma_{\text{total}}$					$\Gamma_{75}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<5.3	90	35 ATHAR	06	CLE3 $\Upsilon(1S) \rightarrow \gamma \pi^+ \pi^-$	

35 Assuming  $B(f_4(2050) \rightarrow \pi\pi) = 0.17$ .

$\Gamma(\gamma f_0(2200) \rightarrow \gamma K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{76}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.0002	90	BARU	89	MD1 $\Upsilon(1S) \rightarrow \gamma K^+ K^-$	

$\Gamma(\gamma f_J(2220) \rightarrow \gamma K^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{77}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 8	90	ATHAR	06	CLE3 $\Upsilon(1S) \rightarrow \gamma K^+ K^-$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 160	90	MASEK	02	CLEO $\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
< 150	90	FULTON	90B	CLEO $\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
< 290	90	ALBRECHT	89	ARG $\Upsilon(1S) \rightarrow \gamma K^+ K^-$	
<2000	90	BARU	89	MD1 $\Upsilon(1S) \rightarrow \gamma K^+ K^-$	

$\Gamma(\gamma f_J(2220) \rightarrow \gamma \pi^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{78}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 6	90	ATHAR	06	CLE3 $\Upsilon(1S) \rightarrow \gamma \pi^+ \pi^-$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<120	90	MASEK	02	CLEO $\Upsilon(1S) \rightarrow \gamma \pi^+ \pi^-$	
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$\Gamma(\gamma f_J(2220) \rightarrow \gamma \rho \bar{\rho})/\Gamma_{\text{total}}$					$\Gamma_{79}/\Gamma$
VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 11	90	ATHAR	06	CLE3 $\Upsilon(1S) \rightarrow \gamma \rho \bar{\rho}$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<160	90	MASEK	02	CLEO $\Upsilon(1S) \rightarrow \gamma \rho \bar{\rho}$	
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$\Gamma(\gamma \eta(2225) \rightarrow \gamma \phi \phi)/\Gamma_{\text{total}}$					$\Gamma_{80}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.003	90	BARU	89	MD1 $\Upsilon(1S) \rightarrow \gamma K^+ K^- K^+ K^-$	

$\Gamma(\gamma \eta_c(1S))/\Gamma_{\text{total}}$					$\Gamma_{81}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<5.7	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma \chi_{c0})/\Gamma_{\text{total}}$					$\Gamma_{82}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<6.5	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma \chi_{c1})/\Gamma_{\text{total}}$					$\Gamma_{83}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<2.3	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma \chi_{c2})/\Gamma_{\text{total}}$					$\Gamma_{84}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<7.6	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma X(3872) \rightarrow \pi^+ \pi^- J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{85}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.6	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma X(3872) \rightarrow \pi^+ \pi^- \pi^0 J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{86}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<2.8	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma X(3915) \rightarrow \omega J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{87}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<3.0	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma X(4140) \rightarrow \phi J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{88}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<2.2	90	SHEN	10A	BELL $\Upsilon(1S) \rightarrow \gamma X$	

$\Gamma(\gamma X)/\Gamma_{\text{total}}$ ( $X = \text{scalar with } m < 8.0 \text{ GeV}$ )					$\Gamma_{89}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

< 4.5	90	36 DEL-AMO-SA..11J	BABR	$e^+e^- \rightarrow \gamma + X$	
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<30	90	37 BALEST	95	CLEO $e^+e^- \rightarrow \gamma + X$	
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36 For a noninteracting scalar  $X$  with mass  $m < 8.0 \text{ GeV}$ .

37 For a noninteracting pseudoscalar  $X$  with mass  $< 7.2 \text{ GeV}$ .

$\Gamma(\gamma X \bar{X}(m_X < 3.1 \text{ GeV}))/\Gamma_{\text{total}}$ ( $X \bar{X} = \text{vectors with } m < 3.1 \text{ GeV}$ )					$\Gamma_{90}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<1	90	38 BALEST	95	CLEO $e^+e^- \rightarrow \gamma + X \bar{X}$	
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38 For a noninteracting vector  $X$  with mass  $< 3.1 \text{ GeV}$ .

$\Gamma(\gamma X \bar{X}(m_X < 4.5 \text{ GeV}))/\Gamma_{\text{total}}$ $X$ and $\bar{X} = \text{zero spin with } m < 4.5 \text{ GeV}$					$\Gamma_{91}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<24	90	39 DEL-AMO-SA..11J	BABR	$e^+e^- \rightarrow \gamma + X \bar{X}$	
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39 For a noninteracting scalar  $X$  with mass  $m < 4.5 \text{ GeV}$ .

$\Gamma(\gamma X \rightarrow \gamma + \geq 4 \text{ prongs})/\Gamma_{\text{total}}$ ( $1.5 \text{ GeV} < m_X < 5.0 \text{ GeV}$ )					$\Gamma_{92}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<1.78	95	ROSNER	07A	CLEO $e^+e^- \rightarrow \gamma X$	
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$\Gamma(\gamma a_1^0 \rightarrow \gamma \mu^+ \mu^-)/\Gamma_{\text{total}}$ ( $201 < M(\mu^+ \mu^-) < 3565 \text{ MeV}$ )					$\Gamma_{93}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<9	90	40 LOVE	08	CLEO $e^+e^- \rightarrow \gamma a_1^0 \rightarrow \gamma \mu^+ \mu^-$	
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<9.7	90	41 LEES	13c	BABR $e^+e^- \rightarrow \gamma a_1^0 \rightarrow \gamma \mu^+ \mu^-$	
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40 For a narrow scalar or pseudoscalar  $a_1^0$  with  $201 < M(\mu^+ \mu^-) < 3565 \text{ MeV}$ , excluding  $J/\psi$ . Measured 90% CL limits as a function of  $M(\mu^+ \mu^-)$  range from  $1-9 \times 10^{-6}$ .

41 For a narrow scalar or pseudoscalar  $a_1^0$  with mass in the range  $212-9200 \text{ MeV}$ , excluding  $J/\psi$  and  $\psi(2S)$ . Measured 90% CL limits as a function of  $m_{a_1^0}$  range from  $0.28-9.7 \times 10^{-6}$ .

$\Gamma(\gamma a_1^0 \rightarrow \gamma \tau^+ \tau^-)/\Gamma_{\text{total}}$ ( $2m_\tau < M(\tau^+ \tau^-) < 9.2 \text{ GeV}$ )					$\Gamma_{94}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<130	90	42 LEES	13R	BABR $\Upsilon(2S) \rightarrow \gamma \tau^+ \tau^- \pi^+ \pi^-$	
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• • • We do not use the following data for averages, fits, limits, etc. • • •

< 50	90	43 LOVE	08	CLEO $e^+e^- \rightarrow \gamma a_1^0 \rightarrow \gamma \tau^+ \tau^-$	
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42 For a narrow scalar  $a_1^0$  with  $2m_\tau < M(a_1^0) < 9.2 \text{ GeV}$ , which result in a 90% CL upper limits of  $0.9 \times 10^{-5}$  at  $M(a_1^0) = 2m_\tau$ ,  $\approx 1.5 \times 10^{-5}$  at  $M(a_1^0) = 7.5 \text{ GeV}$ , and  $13 \times 10^{-5}$  at  $M(a_1^0) = 9.2 \text{ GeV}$ .

43 For a narrow scalar or pseudoscalar  $a_1^0$  with  $2m_\tau < M(a_1^0) < 7.5 \text{ GeV}$ , which result in a 90% CL limits ranging from  $1 \times 10^{-5}$  at  $M(a_1^0)=2m_\tau$  to  $5 \times 10^{-5}$  at  $M(a_1^0)=7.5 \text{ GeV}$ .

$\Gamma(\gamma a_1^0 \rightarrow \gamma g g)/\Gamma_{\text{total}}$ ( $0.5 \text{ GeV} < m < 9.0 \text{ GeV}$ )					$\Gamma_{95}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	

<1 $\times 10^{-2}$	90	44 LEES	13L	BABR $\Upsilon(1S) \rightarrow \gamma X$	
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44 For a narrow,  $CP$ -odd pseudoscalar  $a_1^0$  searched for in 26 hadronic decay modes with invariant mass  $0.5 \text{ GeV} < m_X < 9.0 \text{ GeV}$ . Measured 90% CL limit as a function of  $m_X$  range from  $10^{-6}$  to  $10^{-2}$ .

$\Gamma(\gamma a_1^0 \rightarrow \gamma S \bar{S})/\Gamma_{\text{total}}$ ( $0.5 \text{ GeV} < m < 9.0 \text{ GeV}$ )					$\Gamma_{96}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	

<1 $\times 10^{-3}$	90	45 LEES	13L	BABR $\Upsilon(1S) \rightarrow \gamma X$	
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45 For a narrow,  $CP$ -odd pseudoscalar  $a_1^0$  searched for in 14 hadronic decay modes with invariant mass  $1.5 \text{ GeV} < m_X < 9.0 \text{ GeV}$ . Measured 90% CL limit as a function of  $m_X$  range from  $10^{-5}$  to  $10^{-3}$ .

### LEPTON FAMILY NUMBER (LF) VIOLATING MODES

$\Gamma(\mu^\pm \tau^\mp)/\Gamma_{\text{total}}$					$\Gamma_{97}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

<6.0	95	LOVE	08A	CLEO $e^+e^- \rightarrow \mu^\pm \tau^\mp$	
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### OTHER DECAYS

$\Gamma(\text{invisible})/\Gamma_{\text{total}}$					$\Gamma_{98}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	

< 3.0	90	AUBERT	09AX	BABR $\Upsilon(3S) \rightarrow \pi^+ \pi^- \Upsilon(1S)$	
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<39	90	RUBIN	07	CLEO $\Upsilon(2S) \rightarrow \pi^+ \pi^- \Upsilon(1S)$	
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<25	90	TAJIMA	07	BELL $\Upsilon(3S) \rightarrow \pi^+ \pi^- \Upsilon(1S)$	
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# Meson Particle Listings

## $\chi_{b0}(1P)$ , $\chi_{b1}(1P)$

### $\Gamma(2\pi^+2\pi^-K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_7/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 0.6 \pm 0.1</math></b>	7	10 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^-$
10 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ = $(4 \pm 2 \pm 1) \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = (3.8 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

### $\Gamma(2\pi^+2\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_8/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.7</math></b>	90	11 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^-\pi^0$
11 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 2\pi^+ 2\pi^- K^+ K^-\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $10 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(2\pi^+2\pi^-K^+K^-2\pi^0)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5</math></b>	90	12 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- 2\pi^0$
12 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $20 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(3\pi^+2\pi^-K^-K_S^0\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.6</math></b>	90	13 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 3\pi^+ 2\pi^- K^- K_S^0\pi^0$
13 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 3\pi^+ 2\pi^- K^- K_S^0\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $6 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.8</math></b>	90	14 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 3\pi^+ 3\pi^-$
14 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 3\pi^+ 3\pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $3 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-2\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;6</math></b>	90	15 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 3\pi^+ 3\pi^- 2\pi^0$
15 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 3\pi^+ 3\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $22 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_{13}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.4 \pm 1.2 \pm 0.2</math></b>	9	16 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^-$
16 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 3\pi^+ 3\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ = $(9 \pm 4 \pm 2) \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = (3.8 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

### $\Gamma(3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{14}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;10</math></b>	90	17 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^-\pi^0$
17 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 3\pi^+ 3\pi^- K^+ K^-\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $37 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(4\pi^+4\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{15}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.8</math></b>	90	18 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 4\pi^+ 4\pi^-$
18 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 4\pi^+ 4\pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $3 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(4\pi^+4\pi^-2\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{16}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;21</math></b>	90	19 ASNER	08A CLEO	$\Upsilon(2S) \rightarrow \gamma 4\pi^+ 4\pi^- 2\pi^0$
19 ASNER 08A reports $[\Gamma(\chi_{b0}(1P) \rightarrow 4\pi^+ 4\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ < $77 \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = 3.8 \times 10^{-2}$ .				

### $\Gamma(J/\psi J/\psi)/\Gamma_{\text{total}}$ $\Gamma_{17}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;7</math></b>	90	20 SHEN	12 BELL	$\Upsilon(2S) \rightarrow \gamma \psi X$
20 SHEN 12 reports < $7.1 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b0}(1P) \rightarrow J/\psi J/\psi)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ assuming $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = (3.8 \pm 0.4) \times 10^{-2}$ .				

### $\Gamma(J/\psi\psi(2S))/\Gamma_{\text{total}}$ $\Gamma_{18}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;12</math></b>	90	21 SHEN	12 BELL	$\Upsilon(2S) \rightarrow \gamma \psi X$

21 SHEN 12 reports <  $12 \times 10^{-5}$  from a measurement of  $[\Gamma(\chi_{b0}(1P) \rightarrow J/\psi\psi(2S))/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$  assuming  $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = (3.8 \pm 0.4) \times 10^{-2}$ .

### $\Gamma(\psi(2S)\psi(2S))/\Gamma_{\text{total}}$ $\Gamma_{19}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.1</math></b>	90	22 SHEN	12 BELL	$\Upsilon(2S) \rightarrow \gamma \psi X$
22 SHEN 12 reports < $3.1 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b0}(1P) \rightarrow \psi(2S)\psi(2S))/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))]$ assuming $\text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) = (3.8 \pm 0.4) \times 10^{-2}$ .				

## $\chi_{b0}(1P)$ CROSS-PARTICLE BRANCHING RATIOS

### $\Gamma(\chi_{b0}(1P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.7 \times 10^{-3}</math></b>	90	23 LEES	11J BABR	$\Upsilon(2S) \rightarrow X \gamma$
23 LEES 11J quotes a central value of $\Gamma(\chi_{b0}(1P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))/\Gamma_{\text{total}} = (8.3 \pm 5.6^{+3.7}_{-2.6}) \times 10^{-4}$ and derives a 90% CL upper limit of $\Gamma(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P))/\Gamma_{\text{total}} < 4.6\%$ using $\text{B}(\Upsilon(4S) \rightarrow \gamma \chi_{b0}(1P)) = (3.8 \pm 0.4)\%$ .				

### $\text{B}(\chi_{b0}(1P) \rightarrow \gamma \Upsilon(1S)) \times \text{B}(\Upsilon(2S) \rightarrow \gamma \chi_{b0}(1P)) \times \text{B}(\Upsilon(1S) \rightarrow \ell^+ \ell^-)$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.63 \pm 0.24 \pm 0.15</math></b>	87	KORNICER	11 CLEO	$e^+ e^- \rightarrow \gamma \gamma \ell^+ \ell^-$

## $\chi_{b0}(1P)$ REFERENCES

SHEN	12	PR D85 071102	C.P. Shen <i>et al.</i>	(BELLE Collab.)
KORNICER	11	PR D83 054003	M. Kornicer <i>et al.</i>	(CLEO Collab.)
LEES	11J	PR D84 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ASNER	08A	PR D78 091103	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	08	PR D78 092007	R.A. Briere <i>et al.</i>	(CLEO Collab.)
ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
EDWARDS	99	PR D59 032003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
WALK	86	PR D34 2611	W.S. Walk <i>et al.</i>	(Crystal Ball Collab.)
ALBRECHT	85E	PL 160B 331	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
NERNST	85	PRL 54 2195	R. Nernst <i>et al.</i>	(Crystal Ball Collab.)
HAAS	84	PRL 52 799	J. Haas <i>et al.</i>	(CLEO Collab.)
KLOPFEN...	83	PRL 51 160	C. Klopfenstein <i>et al.</i>	(CUSB Collab.)
PAUSS	83	PL 130B 439	F. Pauss <i>et al.</i>	(MPIM, COLU, CORN, LSU+)

## $\chi_{b1}(1P)$ $I^G(J^{PC}) = 0^+(1^+ +)$ $J$ needs confirmation.

Observed in radiative decay of the  $\Upsilon(2S)$ , therefore  $C = +$ . Branching ratio requires E1 transition, M1 is strongly disfavored, therefore  $P = +$ ,  $J = 1$  from SKWARNICKI 87.

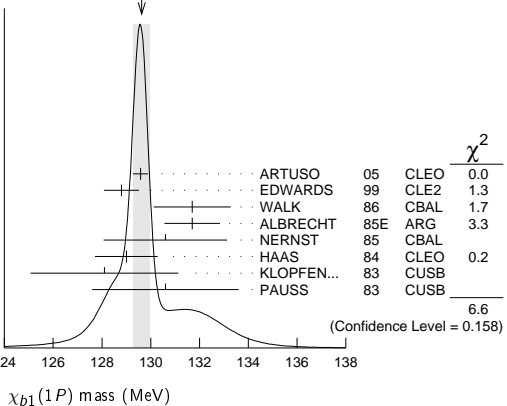
## $\chi_{b1}(1P)$ MASS

VALUE (MeV)	DOCUMENT ID
<b><math>9892.78 \pm 0.26 \pm 0.31</math> OUR EVALUATION</b>	From average $\gamma$ energy below, using $\Upsilon(2S)$ mass = 10023.26 $\pm$ 0.31 MeV

## $\gamma$ ENERGY IN $\Upsilon(2S)$ DECAY

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>129.63 \pm 0.33</math> OUR AVERAGE</b>	Error includes scale factor of 1.3.		See the ideogram below.
129.58 $\pm$ 0.09 $\pm$ 0.29	ARTUSO	05 CLEO	$\Upsilon(2S) \rightarrow \gamma X$
128.8 $\pm$ 0.4 $\pm$ 0.6	EDWARDS	99 CLE2	$\Upsilon(2S) \rightarrow \gamma \chi(1P)$
131.7 $\pm$ 0.9 $\pm$ 1.3	WALK	86 CBAL	$\Upsilon(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$
131.7 $\pm$ 0.3 $\pm$ 1.1	ALBRECHT	85E ARG	$\Upsilon(2S) \rightarrow \text{conv.} \gamma X$
130.6 $\pm$ 0.8 $\pm$ 2.4	NERNST	85 CBAL	$\Upsilon(2S) \rightarrow \gamma X$
129 $\pm$ 0.8 $\pm$ 1	HAAS	84 CLEO	$\Upsilon(2S) \rightarrow \text{conv.} \gamma X$
128.1 $\pm$ 0.4 $\pm$ 3.0	KLOPFEN...	83 CUSB	$\Upsilon(2S) \rightarrow \gamma X$
130.6 $\pm$ 3.0	PAUSS	83 CUSB	$\Upsilon(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$

WEIGHTED AVERAGE  
129.63 $\pm$ 0.33 (Error scaled by 1.3)



$\chi_{b1}(1P)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \quad \gamma \mathcal{T}(1S)$	$(33.9 \pm 2.2) \%$	
$\Gamma_2 \quad D^0 X$	$(12.6 \pm 2.2) \%$	
$\Gamma_3 \quad \pi^+ \pi^- K^+ K^- \pi^0$	$(2.0 \pm 0.6) \times 10^{-4}$	
$\Gamma_4 \quad 2\pi^+ \pi^- K^- K_S^0$	$(1.3 \pm 0.5) \times 10^{-4}$	
$\Gamma_5 \quad 2\pi^+ \pi^- K^- K_S^0 2\pi^0$	$< 6 \times 10^{-4}$	90%
$\Gamma_6 \quad 2\pi^+ 2\pi^- 2\pi^0$	$(8.0 \pm 2.5) \times 10^{-4}$	
$\Gamma_7 \quad 2\pi^+ 2\pi^- K^+ K^-$	$(1.5 \pm 0.5) \times 10^{-4}$	
$\Gamma_8 \quad 2\pi^+ 2\pi^- K^+ K^- \pi^0$	$(3.5 \pm 1.2) \times 10^{-4}$	
$\Gamma_9 \quad 2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	$(8.6 \pm 3.2) \times 10^{-4}$	
$\Gamma_{10} \quad 3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	$(9.3 \pm 3.3) \times 10^{-4}$	
$\Gamma_{11} \quad 3\pi^+ 3\pi^-$	$(1.9 \pm 0.6) \times 10^{-4}$	
$\Gamma_{12} \quad 3\pi^+ 3\pi^- 2\pi^0$	$(1.7 \pm 0.5) \times 10^{-3}$	
$\Gamma_{13} \quad 3\pi^+ 3\pi^- K^+ K^-$	$(2.6 \pm 0.8) \times 10^{-4}$	
$\Gamma_{14} \quad 3\pi^+ 3\pi^- K^+ K^- \pi^0$	$(7.5 \pm 2.6) \times 10^{-4}$	
$\Gamma_{15} \quad 4\pi^+ 4\pi^-$	$(2.6 \pm 0.9) \times 10^{-4}$	
$\Gamma_{16} \quad 4\pi^+ 4\pi^- 2\pi^0$	$(1.4 \pm 0.6) \times 10^{-3}$	
$\Gamma_{17} \quad J/\psi J/\psi$	$< 2.7 \times 10^{-5}$	90%
$\Gamma_{18} \quad J/\psi \psi(2S)$	$< 1.7 \times 10^{-5}$	90%
$\Gamma_{19} \quad \psi(2S) \psi(2S)$	$< 6 \times 10^{-5}$	90%

 $\chi_{b1}(1P)$  BRANCHING RATIOS

$\Gamma(\gamma \mathcal{T}(1S))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>0.339 \pm 0.022</math> OUR AVERAGE</b>	
$0.331 \pm 0.018 \pm 0.017$ 3222 <sup>1,2</sup> KORNICER 11 CLEO $e^+ e^- \rightarrow \gamma \gamma \ell^+ \ell^-$	
$0.350 \pm 0.023 \pm 0.018$ 13k <sup>3</sup> LEES 11j BABR $\mathcal{T}(2S) \rightarrow X \gamma$	
$0.32 \pm 0.06 \pm 0.07$ WALK 86 CBAL $\mathcal{T}(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$	
$0.47 \pm 0.18$ KLOPFEN... 83 CUSB $\mathcal{T}(2S) \rightarrow \gamma \gamma \ell^+ \ell^-$	

<sup>1</sup> Assuming  $B(\mathcal{T}(1S) \rightarrow \ell^+ \ell^-) = (2.48 \pm 0.05)\%$ .

<sup>2</sup> KORNICER 11 reports  $[\Gamma(\chi_{b1}(1P) \rightarrow \gamma \mathcal{T}(1S))/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (22.8 \pm 0.4 \pm 1.2) \times 10^{-3}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> LEES 11j reports  $[\Gamma(\chi_{b1}(1P) \rightarrow \gamma \mathcal{T}(1S))/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (24.1 \pm 0.6 \pm 1.5) \times 10^{-3}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D^0 X)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
<b>VALUE (units <math>10^{-2}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>12.6 \pm 1.9 \pm 1.1</math></b> 2310 <sup>4</sup> BRIERE 08 CLEO $\mathcal{T}(2S) \rightarrow \gamma D^0 X$	

<sup>4</sup> For  $p_{D^0} > 2.5 \text{ GeV}/c$ .

$\Gamma(\pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>2.0 \pm 0.6 \pm 0.1</math></b> 18 <sup>5</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma \pi^+ \pi^- K^+ K^- \pi^0$	

<sup>5</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow \pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (14 \pm 3 \pm 3) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>1.3 \pm 0.5 \pm 0.1</math></b> 11 <sup>6</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 2\pi^+ \pi^- K^- K_S^0$	

<sup>6</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (9 \pm 3 \pm 2) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>CL%</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>&lt; 6</math></b> 90 <sup>7</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 2\pi^+ \pi^- K^- 2\pi^0$	

<sup>7</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $< 42 \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = 6.9 \times 10^{-2}$ .

$\Gamma(2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>8.0 \pm 2.4 \pm 0.4</math></b> 46 <sup>8</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 2\pi^+ 2\pi^- 2\pi^0$	

<sup>8</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (55 \pm 9 \pm 14) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>1.5 \pm 0.5 \pm 0.1</math></b> 18 <sup>9</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^-$	

<sup>9</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (10 \pm 3 \pm 2) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2\pi^+ 2\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>3.5 \pm 1.2 \pm 0.2</math></b> 22 <sup>10</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- \pi^0$	

<sup>10</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (24 \pm 6 \pm 6) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>8.6 \pm 3.2 \pm 0.4</math></b> 26 <sup>11</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	

<sup>11</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (59 \pm 14 \pm 17) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+ 2\pi^- K^- K_S^0 \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>9.3 \pm 3.3 \pm 0.5</math></b> 21 <sup>12</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	

<sup>12</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 3\pi^+ 2\pi^- K^- K_S^0 \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (64 \pm 16 \pm 16) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+ 3\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>1.9 \pm 0.6 \pm 0.1</math></b> 25 <sup>13</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 3\pi^+ 3\pi^-$	

<sup>13</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 3\pi^+ 3\pi^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (13 \pm 3 \pm 3) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+ 3\pi^- 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>17 \pm 5 \pm 1</math></b> 56 <sup>14</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 3\pi^+ 3\pi^- 2\pi^0$	

<sup>14</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 3\pi^+ 3\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (119 \pm 18 \pm 32) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+ 3\pi^- K^+ K^-)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>2.6 \pm 0.8 \pm 0.1</math></b> 21 <sup>15</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^-$	

<sup>15</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 3\pi^+ 3\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (18 \pm 4 \pm 4) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+ 3\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{14}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>7.5 \pm 2.6 \pm 0.4</math></b> 28 <sup>16</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^- \pi^0$	

<sup>16</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 3\pi^+ 3\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (52 \pm 11 \pm 14) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(4\pi^+ 4\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{15}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>2.6 \pm 0.9 \pm 0.1</math></b> 24 <sup>17</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 4\pi^+ 4\pi^-$	

<sup>17</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 4\pi^+ 4\pi^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (18 \pm 4 \pm 5) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(4\pi^+ 4\pi^- 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{16}/\Gamma$
<b>VALUE (units <math>10^{-4}</math>)</b> <b>EVTS</b> <b>DOCUMENT ID</b> <b>TECN</b> <b>COMMENT</b>	
<b><math>14 \pm 5 \pm 1</math></b> 26 <sup>18</sup> ASNER 08A CLEO $\mathcal{T}(2S) \rightarrow \gamma 4\pi^+ 4\pi^- 2\pi^0$	

<sup>18</sup> ASNER 08A reports  $[\Gamma(\chi_{b1}(1P) \rightarrow 4\pi^+ 4\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P))]$   $= (96 \pm 24 \pm 29) \times 10^{-6}$  which we divide by our best value  $B(\mathcal{T}(2S) \rightarrow \gamma \chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

## $\chi_{b1}(1P)$ , $h_b(1P)$ , $\chi_{b2}(1P)$

$\Gamma(J/\psi J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<2.7	90	19 SHEN	12 BELL	$\Upsilon(2S) \rightarrow \gamma\psi X$	
19 SHEN 12 reports $< 2.7 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b1}(1P) \rightarrow J/\psi J/\psi)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P))]$ assuming $\text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ .					

$\Gamma(J/\psi\psi(2S))/\Gamma_{\text{total}}$				$\Gamma_{18}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<1.7	90	<sup>20</sup> SHEN	12 BELL	$\Upsilon(2S) \rightarrow \gamma\psi X$	
<sup>20</sup> SHEN 12 reports $< 1.7 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b1}(1P) \rightarrow J/\psi\psi(2S))/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P))]$ assuming $\text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ .					

$\Gamma(\psi(2S)\psi(2S))/\Gamma_{\text{total}}$				$\Gamma_{19}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<6	90	<sup>21</sup> SHEN	12 BELL	$\Upsilon(2S) \rightarrow \gamma\psi X$	
<sup>21</sup> SHEN 12 reports $< 6.2 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b1}(1P) \rightarrow \psi(2S)\psi(2S))/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P))]$ assuming $\text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P)) = (6.9 \pm 0.4) \times 10^{-2}$ .					

### $\chi_{b1}(1P)$ Cross-Particle Branching Ratios

$\Gamma(\chi_{b1}(1P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P))/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma \times \Gamma_{45}^{\Upsilon(2S)}/\Gamma \Upsilon(2S)$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
24.1±0.6±1.5	13k	LEES	11j BABR	$\Upsilon(2S) \rightarrow X\gamma$	

$\text{B}(\chi_{b1}(1P) \rightarrow \gamma \Upsilon(1S)) \times \text{B}(\Upsilon(2S) \rightarrow \gamma\chi_{b1}(1P)) \times \text{B}(\Upsilon(1S) \rightarrow \ell^+ \ell^-)$					
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
5.65±0.11±0.27	3222	KORNICER	11 CLEO	$e^+e^- \rightarrow \gamma\gamma\ell^+\ell^-$	

$\text{B}(\chi_{b1}(1P) \rightarrow \gamma \Upsilon(1S)) \times \text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(1P)) \times \text{B}(\Upsilon(1S) \rightarrow \ell^+ \ell^-)$					
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
1.33±0.30±0.23	50	KORNICER	11 CLEO	$e^+e^- \rightarrow \gamma\gamma\ell^+\ell^-$	

$\text{B}(\chi_{b2}(1P) \rightarrow \rho X + \bar{\rho} X)/\text{B}(\chi_{b1}(1P) \rightarrow \rho X + \bar{\rho} X)$			
VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.068±0.010±0.040</b>	BRIERE	07 CLEO	$\Upsilon(2S) \rightarrow \gamma \chi_{bJ}(1P)$

$\text{B}(\chi_{b0}(1P) \rightarrow \rho X + \bar{\rho} X)/\text{B}(\chi_{b1}(1P) \rightarrow \rho X + \bar{\rho} X)$				
VALUE	DOCUMENT ID	TECN	COMMENT	
$1.11 \pm 0.15 \pm 0.20$	BRIERE	07 CLEO	$\Upsilon(2S) \rightarrow \gamma \chi_{bJ}(1P)$	

### $\chi_{b1}(1P)$ REFERENCES

SHEN	12	PR D85 071102	C.P. Shen <i>et al.</i>	(BELLE Collab.)
KORNICER	11	PR D83 054003	M. Kornicer <i>et al.</i>	(CLEO Collab.)
LEES	11j	PR D84 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ASNER	08a	PR D78 091103	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	08a	PR D78 092007	R.A. Briere <i>et al.</i>	(CLEO Collab.)
BRIERE	07	PR D76 012005	R.A. Briere <i>et al.</i>	(CLEO Collab.)
ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
EDWARDS	99	PR D59 032003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
SKWARNICKI	87	PRL 58 972	T. Skwarnicki <i>et al.</i>	(Crystal Ball Collab.)
WALK	86	PR D34 2611	W.S. Walk <i>et al.</i>	(Crystal Ball Collab.)
ALBRECHT	85E	PL 160B 331	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
NERNST	85	PRL 54 2195	R. Nernst <i>et al.</i>	(Crystal Ball Collab.)
HAAS	84	PRL 52 799	J. Haas <i>et al.</i>	(CLEO Collab.)
KLOPFEN...	83	PRL 51 160	C. Klopfenstein <i>et al.</i>	(CUSB Collab.)
PAUSS	83	PL 130B 439	F. Pauss <i>et al.</i>	(MPIM, COLU, CORN, LSU+)

## $h_b(1P)$

$$I^G(J^{PC}) = ?^?(1^{+-})$$

Quantum numbers are quark model predictions,  $C = -$  established by  $\eta_b\gamma$  decay.

### $h_b(1P)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9899.3±0.8 OUR AVERAGE</b>				
9899.3±0.4±1.0	112k	TAMPONI	15 BELL	$e^+e^- \rightarrow \gamma\eta + \text{hadrons}$
9899.1±0.4±1.0	70k	MIZUK	12 BELL	$e^+e^- \rightarrow \pi^+\pi^- \text{hadrons}$
9902 ±4 ±2	10.8k	LEES	11k BABR	$\Upsilon(3S) \rightarrow \eta_b \gamma \pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
9898.2+1.1+1.0 -1.0-1.1	50.0k	<sup>1</sup> ADACHI	12 BELL	10.86 $e^+e^- \rightarrow \pi^+\pi^- \text{MM}$
<sup>1</sup> Superseded by MIZUK 12.				

### $h_b(1P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \eta_b(1S)\gamma$	(52+6-5)%

### $h_b(1P)$ BRANCHING RATIOS

$\Gamma(\eta_b(1S)\gamma)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
$52 \pm 5$	OUR AVERAGE				
$56 \pm 8 \pm 4$	33.1k	<sup>1</sup> TAMPONI	15 BELL	$e^+e^- \rightarrow \gamma\eta + \text{hadrons}$	
$49.2 \pm 5.7 \pm 5.6_{-3.3}$	24k	MIZUK	12 BELL	$e^+e^- \rightarrow (\gamma)\pi^+\pi^- \text{ hadrons}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen	10.8k	LEES	11k BABR	$\Upsilon(3S) \rightarrow \eta_b\gamma\pi^0$	
<sup>1</sup> Using $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20)\%$ .					

### $h_b(1P)$ REFERENCES

TAMPONI	15	PRL 115 142001	U. Tamponi <i>et al.</i>	(BELLE Collab.)
ADACHI	12	PRL 108 032001	I. Adachi <i>et al.</i>	(BELLE Collab.)
MIZUK	12	PRL 109 232002	R. Mizuk <i>et al.</i>	(BELLE Collab.)
LEES	11k	PR D84 091101	J.P. Lees <i>et al.</i>	(BABAR Collab.)

## $\chi_{b2}(1P)$

$$I^G(J^{PC}) = 0^+(2^{++})$$

$J$  needs confirmation.

Observed in radiative decay of the  $\Upsilon(2S)$ , therefore  $C = +$ . Branching ratio requires E1 transition, M1 is strongly disfavored, therefore  $P = +$ .  $J = 2$  from SKWARNICKI 87.

### $\chi_{b2}(1P)$ MASS

VALUE (MeV)	DOCUMENT ID
9912.21±0.26±0.31 OUR EVALUATION	From average $\gamma$ energy below, using $\Upsilon(2S)$ mass = 10023.26 ± 0.31 MeV

### $m_{\chi_{b2}(1P)} - m_{\chi_{b1}(1P)}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
19.81±0.65±0.20	1 AAIJ	14Bg LHCb	$p\bar{p} \rightarrow \gamma\mu^+\mu^-X$
1 From the $\chi_{b2}(1P) \rightarrow \Upsilon(1S)\gamma$ transition.			

### $\gamma$ ENERGY IN $\Upsilon(2S)$ DECAY

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
110.44±0.29 OUR AVERAGE	Error includes scale factor of 1.1.		
110.58±0.08±0.30	ARTUSO	05 CLEO	$\Upsilon(2S) \rightarrow \gamma X$
110.8 ±0.3 ±0.6	EDWARDS	99 CLE2	$\Upsilon(2S) \rightarrow \gamma\chi(1P)$
107.0 ±1.1 ±1.3	WALK	86 CBAL	$\Upsilon(2S) \rightarrow \gamma\gamma\ell^+\ell^-$
110.6 ±0.3 ±0.9	ALBRECHT	85E ARG	$\Upsilon(2S) \rightarrow \text{conv.}\gamma X$
110.4 ±0.8 ±2.2	NERNST	85 CBAL	$\Upsilon(2S) \rightarrow \gamma X$
109.5 ±0.7 ±1.0	HAAS	84 CLEO	$\Upsilon(2S) \rightarrow \text{conv.}\gamma X$
108.2 ±0.3 ±2.0	KLOPFEN...	83 CUSB	$\Upsilon(2S) \rightarrow \gamma X$
108.8 ±4.0	PAUSS	83 CUSB	$\Upsilon(2S) \rightarrow \gamma\gamma\ell^+\ell^-$

### $\chi_{b2}(1P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \quad \gamma \Upsilon(1S)$	(19.1±1.2)%	
$\Gamma_2 \quad D^0 X$	< 7.9%	90%
$\Gamma_3 \quad \pi^+\pi^- K^+ K^- \pi^0$	( 8 ±5 ) × 10 <sup>-5</sup>	
$\Gamma_4 \quad 2\pi^+\pi^- K^- K_S^0$	< 1.0 × 10 <sup>-4</sup>	90%
$\Gamma_5 \quad 2\pi^+\pi^- K^- K_S^0 2\pi^0$	( 5.3±2.4 ) × 10 <sup>-4</sup>	
$\Gamma_6 \quad 2\pi^+2\pi^- 2\pi^0$	( 3.5±1.4 ) × 10 <sup>-4</sup>	
$\Gamma_7 \quad 2\pi^+2\pi^- K^+ K^-$	( 1.1±0.4 ) × 10 <sup>-4</sup>	
$\Gamma_8 \quad 2\pi^+2\pi^- K^+ K^- \pi^0$	( 2.1±0.9 ) × 10 <sup>-4</sup>	
$\Gamma_9 \quad 2\pi^+2\pi^- K^+ K^- 2\pi^0$	( 3.9±1.8 ) × 10 <sup>-4</sup>	
$\Gamma_{10} \quad 3\pi^+2\pi^- K^- K_S^0 \pi^0$	< 5 × 10 <sup>-4</sup>	90%
$\Gamma_{11} \quad 3\pi^+3\pi^-$	( 7.0±3.1 ) × 10 <sup>-5</sup>	
$\Gamma_{12} \quad 3\pi^+3\pi^- 2\pi^0$	( 1.0±0.4 ) × 10 <sup>-3</sup>	
$\Gamma_{13} \quad 3\pi^+3\pi^- K^+ K^-$	< 8 × 10 <sup>-5</sup>	90%
$\Gamma_{14} \quad 3\pi^+3\pi^- K^+ K^- \pi^0$	( 3.6±1.5 ) × 10 <sup>-4</sup>	
$\Gamma_{15} \quad 4\pi^+4\pi^-$	( 8 ±4 ) × 10 <sup>-5</sup>	
$\Gamma_{16} \quad 4\pi^+4\pi^- 2\pi^0$	( 1.8±0.7 ) × 10 <sup>-3</sup>	
$\Gamma_{17} \quad J/\psi J/\psi$	< 4 × 10 <sup>-5</sup>	90%
$\Gamma_{18} \quad J/\psi\psi(2S)$	< 5 × 10 <sup>-5</sup>	90%
$\Gamma_{19} \quad \psi(2S)\psi(2S)$	< 1.6 × 10 <sup>-5</sup>	90%

$\chi_{b2}(1P)$  BRANCHING RATIOS

$\Gamma(\gamma T(1S))/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.191 \pm 0.012</math> OUR AVERAGE</b>					
$0.186 \pm 0.011 \pm 0.009$	1770	<sup>2,3</sup> KORNICER	11 CLEO	$e^+e^- \rightarrow \gamma\gamma\ell^+\ell^-$	
$0.194^{+0.014}_{-0.017} \pm 0.009$	8k	<sup>4</sup> LEES	11j BABR	$T(2S) \rightarrow X\gamma$	
$0.27 \pm 0.06 \pm 0.06$		WALK	86 CBAL	$T(2S) \rightarrow \gamma\gamma\ell^+\ell^-$	
$0.20 \pm 0.05$		KLOPFEN...	83 CUSB	$T(2S) \rightarrow \gamma\gamma\ell^+\ell^-$	

<sup>2</sup> Assuming  $B(T(1S) \rightarrow \ell^+\ell^-) = (2.48 \pm 0.05)\%$ .

<sup>3</sup> KORNICER 11 reports  $[\Gamma(\chi_{b2}(1P) \rightarrow \gamma T(1S))/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$  =  $(1.33 \pm 0.04 \pm 0.07) \times 10^{-2}$  which we divide by our best value  $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$  =  $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> LEES 11j reports  $[\Gamma(\chi_{b2}(1P) \rightarrow \gamma T(1S))/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$  =  $(13.9 \pm 0.5^{+0.9}_{-1.1}) \times 10^{-3}$  which we divide by our best value  $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$  =  $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(D^0 X)/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;7.9 \times 10^{-2}</math></b>	90	<sup>5,6</sup> BRIERE	08 CLEO	$T(2S) \rightarrow \gamma D^0 X$	
<sup>5</sup> For $p_{D^0} > 2.5$ GeV/c.					
<sup>6</sup> The authors also present their result as $(5.4 \pm 1.9 \pm 0.5) \times 10^{-2}$ .					

$\Gamma(\pi^+\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.84 \pm 0.50 \pm 0.04</math></b>	8	<sup>7</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma\pi^+\pi^-K^+K^-\pi^0$	
<sup>7</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow \pi^+\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(6 \pm 3 \pm 2) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(2\pi^+\pi^-K^-K_S^0)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.0</math></b>	90	<sup>8</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 2\pi^+\pi^-K^-K_S^0$	
<sup>8</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 2\pi^+\pi^-K^-K_S^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $< 7 \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $7.15 \times 10^{-2}$ .					

$\Gamma(2\pi^+\pi^-K^-K_S^0 2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>5.3 \pm 2.4 \pm 0.3</math></b>	11	<sup>9</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 2\pi^+\pi^-K^-2\pi^0$	
<sup>9</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 2\pi^+\pi^-K^-K_S^0 2\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(38 \pm 14 \pm 10) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(2\pi^+2\pi^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.5 \pm 1.4 \pm 0.2</math></b>	19	<sup>10</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 2\pi^+2\pi^-2\pi^0$	
<sup>10</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 2\pi^+2\pi^-2\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(25 \pm 8 \pm 6) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(2\pi^+2\pi^-K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.1 \pm 0.4 \pm 0.1</math></b>	14	<sup>11</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 2\pi^+2\pi^-K^+K^-$	
<sup>11</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 2\pi^+2\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(8 \pm 2 \pm 2) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(2\pi^+2\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.1 \pm 0.9 \pm 0.1</math></b>	13	<sup>12</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 2\pi^+2\pi^-K^+K^-\pi^0$	
<sup>12</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 2\pi^+2\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(15 \pm 5 \pm 4) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(2\pi^+2\pi^-K^+K^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.9 \pm 1.8 \pm 0.2</math></b>	11	<sup>13</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 2\pi^+2\pi^-K^+K^-2\pi^0$	
<sup>13</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 2\pi^+2\pi^-K^+K^-2\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(28 \pm 11 \pm 7) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(3\pi^+2\pi^-K^-K_S^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{10}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;5</math></b>	90	<sup>14</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 3\pi^+2\pi^-K^-K_S^0\pi^0$	
<sup>14</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 3\pi^+2\pi^-K^-K_S^0\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $< 36 \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $7.15 \times 10^{-2}$ .					

$\Gamma(3\pi^+3\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.70 \pm 0.31 \pm 0.03</math></b>	9	<sup>15</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 3\pi^+3\pi^-$	
<sup>15</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 3\pi^+3\pi^-)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(5 \pm 2 \pm 1) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(3\pi^+3\pi^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>10.2 \pm 3.6 \pm 0.5</math></b>	34	<sup>16</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 3\pi^+3\pi^-2\pi^0$	
<sup>16</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 3\pi^+3\pi^-2\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(73 \pm 16 \pm 20) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(3\pi^+3\pi^-K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;0.8</math></b>	90	<sup>17</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 3\pi^+3\pi^-K^+K^-$	
<sup>17</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 3\pi^+3\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $< 6 \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $7.15 \times 10^{-2}$ .					

$\Gamma(3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{14}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.6 \pm 1.5 \pm 0.2</math></b>	14	<sup>18</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 3\pi^+3\pi^-K^+K^-\pi^0$	
<sup>18</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(26 \pm 8 \pm 7) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(4\pi^+4\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.84 \pm 0.40 \pm 0.04</math></b>	7	<sup>19</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 4\pi^+4\pi^-$	
<sup>19</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 4\pi^+4\pi^-)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(6 \pm 2 \pm 2) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(4\pi^+4\pi^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{16}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>18 \pm 7 \pm 1</math></b>	29	<sup>20</sup> ASNER	08A CLEO	$T(2S) \rightarrow \gamma 4\pi^+4\pi^-2\pi^0$	
<sup>20</sup> ASNER 08A reports $[\Gamma(\chi_{b2}(1P) \rightarrow 4\pi^+4\pi^-2\pi^0)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ = $(132 \pm 31 \pm 40) \times 10^{-6}$ which we divide by our best value $B(T(2S) \rightarrow \gamma\chi_{b2}(1P))$ = $(7.15 \pm 0.35) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.					

$\Gamma(J/\psi J/\psi)/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;5</math></b>	90	<sup>21</sup> SHEN	12 BELL	$T(2S) \rightarrow \gamma\psi X$	
<sup>21</sup> SHEN 12 reports $< 4.5 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b2}(1P) \rightarrow J/\psi J/\psi)/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ assuming $B(T(2S) \rightarrow \gamma\chi_{b2}(1P)) = (7.15 \pm 0.35) \times 10^{-2}$ .					

$\Gamma(J/\psi\psi(2S))/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;5</math></b>	90	<sup>22</sup> SHEN	12 BELL	$T(2S) \rightarrow \gamma\psi X$	
<sup>22</sup> SHEN 12 reports $< 4.9 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b2}(1P) \rightarrow J/\psi\psi(2S))/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ assuming $B(T(2S) \rightarrow \gamma\chi_{b2}(1P)) = (7.15 \pm 0.35) \times 10^{-2}$ .					

$\Gamma(\psi(2S)\psi(2S))/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;1.6</math></b>	90	<sup>23</sup> SHEN	12 BELL	$T(2S) \rightarrow \gamma\psi X$	
<sup>23</sup> SHEN 12 reports $< 1.6 \times 10^{-5}$ from a measurement of $[\Gamma(\chi_{b2}(1P) \rightarrow \psi(2S)\psi(2S))/\Gamma_{\text{total}}] \times [B(T(2S) \rightarrow \gamma\chi_{b2}(1P))]$ assuming $B(T(2S) \rightarrow \gamma\chi_{b2}(1P)) = (7.15 \pm 0.35) \times 10^{-2}$ .					

Meson Particle Listings

$\chi_{b2}(1P)$ ,  $\eta_b(2S)$ ,  $\Upsilon(2S)$

$\chi_{b2}(1P)$ Cross-Particle Branching Ratios				
$\Gamma(\chi_{b2}(1P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(2S) \rightarrow \gamma \chi_{b2}(1P))/\Gamma_{\text{total}}$				
$\Gamma_1/\Gamma \times \Gamma_{46}^{\Upsilon(2S)}/\Gamma \Upsilon(2S)$				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>13.9 \pm 0.5^{+0.9}_{-1.1}</math></b>	8k	LEES	11j	BABR $\Upsilon(2S) \rightarrow X \gamma$
$B(\chi_{b2}(1P) \rightarrow \gamma \Upsilon(1S)) \times B(\Upsilon(2S) \rightarrow \gamma \chi_{b2}(1P)) \times B(\Upsilon(1S) \rightarrow \ell^+ \ell^-)$				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.29 \pm 0.09 \pm 0.16</math></b>	1770	KORNICER	11	CLEO $e^+ e^- \rightarrow \gamma \gamma \ell^+ \ell^-$
$B(\chi_{b2}(1P) \rightarrow \gamma \Upsilon(1S)) \times B(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(1P)) \times B(\Upsilon(1S) \rightarrow \ell^+ \ell^-)$				
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.56 \pm 0.40 \pm 0.41</math></b>	126	KORNICER	11	CLEO $e^+ e^- \rightarrow \gamma \gamma \ell^+ \ell^-$

$\chi_{b2}(1P)$ REFERENCES				
AJU	14BG	JHEP 1410 088	R. Aaij <i>et al.</i>	(LHCb Collab.)
SHEN	12	PR D85 071102	C.P. Shen <i>et al.</i>	(BELLE Collab.)
KORNICER	11	PR D83 054003	M. Kornicer <i>et al.</i>	(CLEO Collab.)
LEES	11j	PR D84 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ASNER	08a	PR D78 091103	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	08	PR D78 092007	R.A. Briere <i>et al.</i>	(CLEO Collab.)
ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
EDWARDS	99	PR D59 032003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
SKWARNICKI	87	PRL 58 972	T. Skwarnicki <i>et al.</i>	(Crystal Ball Collab.) J
WALK	86	PR D34 2611	W.S. Walk <i>et al.</i>	(Crystal Ball Collab.)
ALBRECHT	85E	PL 160B 331	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
NERNST	85	PRL 54 2195	R. Nernst <i>et al.</i>	(Crystal Ball Collab.)
HAAS	84	PRL 52 799	J. Haas <i>et al.</i>	(CLEO Collab.)
KLOPFEN...	83	PRL 51 160	C. Klopfenstein <i>et al.</i>	(CUSB Collab.)
PAUSS	83	PL 130B 439	F. Pauss <i>et al.</i>	(MPIM, COLU, CORN, LSU+)

$\eta_b(2S)$	$I^G(J^{PC}) = 0^+(0^-+)$
OMITTED FROM SUMMARY TABLE	
Quantum numbers shown are quark-model predictions.	

$\eta_b(2S)$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>9999.0 \pm 3.5^{+2.8}_{-1.9}</math></b>	26k	<sup>1</sup> MIZUK	12	BELL $e^+ e^- \rightarrow \gamma \pi^+ \pi^- +$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$9974.6 \pm 2.3 \pm 2.1$	$11 \pm 4$	<sup>2,3</sup> DOBBS	12	$\Upsilon(2S) \rightarrow \gamma$ hadrons
<sup>1</sup> Assuming $\Gamma_{\eta_b(2S)} = 4.9$ MeV. Not independent of the corresponding mass difference measurement.				
<sup>2</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.				
<sup>3</sup> Assuming $\Gamma_{\eta_b(2S)} = 5$ MeV. Not independent of the corresponding mass difference measurement.				
$m_{\Upsilon(2S)} - m_{\eta_b(2S)}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>24.3 \pm 3.5^{+2.8}_{-1.9}</math></b>	26k	<sup>4</sup> MIZUK	12	BELL $e^+ e^- \rightarrow \gamma \pi^+ \pi^- +$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$48.7 \pm 2.3 \pm 2.1$	$11 \pm 4$	<sup>5,6</sup> DOBBS	12	$\Upsilon(2S) \rightarrow \gamma$ hadrons
<sup>4</sup> Assuming $\Gamma_{\eta_b(2S)} = 4.9$ MeV. Not independent of the corresponding mass measurement.				
<sup>5</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.				
<sup>6</sup> Assuming $\Gamma_{\eta_b(2S)} = 5$ MeV. Not independent of the corresponding mass measurement.				

$\eta_b(2S)$ WIDTH				
VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;24</b>	90	MIZUK	12	BELL $e^+ e^- \rightarrow \gamma \pi^+ \pi^-$ hadrons

$\eta_b(2S)$  DECAY MODES
 

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ hadrons	seen

$\eta_b(2S)$  BRANCHING RATIOS
 

$\Gamma(\text{hadrons})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	26k	MIZUK	12	BELL $e^+e^- \rightarrow \gamma\pi^+\pi^-$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	<sup>7</sup>	DOBBS	12	$\Upsilon(2S) \rightarrow \gamma$ hadrons
<sup>7</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.				

$\eta_b(2S)$  REFERENCES
 

DOBBS	12	PRL 109 082001	S. Dobbs <i>et al.</i>	
MIZUK	12	PRL 109 232002	R. Mizuk <i>et al.</i>	(BELLE Collab.)

$\Upsilon(2S)$	$I^G(J^{PC}) = 0^-(1^{--})$			
$T(2S) \text{ MASS}$				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b><math>10023.26 \pm 0.31</math> OUR AVERAGE</b>				
$10023.5 \pm 0.5$	<sup>1</sup> ARTA MONOV 00	MD1	$e^+e^- \rightarrow$ hadrons	
$10023.1 \pm 0.4$	BARBER 84	REDE	$e^+e^- \rightarrow$ hadrons	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$10023.6 \pm 0.5$	<sup>2,3</sup> BARU	86B	REDE	$e^+e^- \rightarrow$ hadrons
<sup>1</sup> Reanalysis of BARU 86B using new electron mass (COHEN 87).				
<sup>2</sup> Reanalysis of ARTAMONOV 84.				
<sup>3</sup> Superseded by ARTAMONOV 00.				

$m_{\Upsilon(3S)} - m_{\Upsilon(2S)}$				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b><math>331.50 \pm 0.02 \pm 0.13</math></b>	LEES	11c	BABR	$e^+ e^- \rightarrow \pi^+ \pi^- X$

$\Upsilon(2S)$ WIDTH	
VALUE (keV)	DOCUMENT ID
<b><math>31.98 \pm 2.63</math> OUR EVALUATION</b>	See the Note on “Width Determinations of the $\Upsilon$ States”

$\Upsilon(2S)$ DECAY MODES		Scale factor/ Confidence level	
Mode	Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1 \Upsilon(1S) \pi^+ \pi^-$	$(17.85 \pm 0.26) \%$		
$\Gamma_2 \Upsilon(1S) \pi^0 \pi^0$	$( 8.6 \pm 0.4 ) \%$		
$\Gamma_3 \tau^+ \tau^-$	$( 2.00 \pm 0.21 ) \%$		
$\Gamma_4 \mu^+ \mu^-$	$( 1.93 \pm 0.17 ) \%$	S=2.2	
$\Gamma_5 e^+ e^-$	$( 1.91 \pm 0.16 ) \%$		
$\Gamma_6 \Upsilon(1S) \pi^0$	$< 4$	$\times 10^{-5}$	CL=90%
$\Gamma_7 \Upsilon(1S) \eta$	$( 2.9 \pm 0.4 ) \times 10^{-4}$	$\times 10^{-4}$	S=2.0
$\Gamma_8 J/\psi(1S)$ anything	$< 6$	$\times 10^{-3}$	CL=90%
$\Gamma_9 J/\psi(1S) \eta_c$	$< 5.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{10} J/\psi(1S) \chi_{c0}$	$< 3.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{11} J/\psi(1S) \chi_{c1}$	$< 1.2$	$\times 10^{-6}$	CL=90%
$\Gamma_{12} J/\psi(1S) \chi_{c2}$	$< 2.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{13} J/\psi(1S) \eta_c(2S)$	$< 2.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{14} J/\psi(1S) X(3940)$	$< 2.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{15} J/\psi(1S) X(4160)$	$< 2.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{16} \psi(2S) \eta_c$	$< 5.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{17} \psi(2S) \chi_{c0}$	$< 4.7$	$\times 10^{-6}$	CL=90%
$\Gamma_{18} \psi(2S) \chi_{c1}$	$< 2.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{19} \psi(2S) \chi_{c2}$	$< 1.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{20} \psi(2S) \eta_c(2S)$	$< 3.3$	$\times 10^{-6}$	CL=90%
$\Gamma_{21} \psi(2S) X(3940)$	$< 3.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{22} \psi(2S) X(4160)$	$< 3.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{23} \overline{2}H$ anything	$( 2.78 \pm 0.30_{-0.26} ) \times 10^{-5}$	$\times 10^{-5}$	S=1.2
$\Gamma_{24}$ hadrons	$(94 \pm 11) \%$		
$\Gamma_{25} g g g$	$(58.8 \pm 1.2) \%$		
$\Gamma_{26} \gamma g g$	$( 1.87 \pm 0.28 ) \%$		
$\Gamma_{27} \phi K^+ K^-$	$( 1.6 \pm 0.4 ) \times 10^{-6}$		
$\Gamma_{28} \omega \pi^+ \pi^-$	$< 2.58$	$\times 10^{-6}$	CL=90%
$\Gamma_{29} K^*(892)^0 K^- \pi^+ + \text{c.c.}$	$( 2.3 \pm 0.7 ) \times 10^{-6}$	$\times 10^{-6}$	
$\Gamma_{30} \phi f'_2(1525)$	$< 1.33$	$\times 10^{-6}$	CL=90%
$\Gamma_{31} \omega f_2(1270)$	$< 5.7$	$\times 10^{-7}$	CL=90%
$\Gamma_{32} \rho(770) a_2(1320)$	$< 8.8$	$\times 10^{-7}$	CL=90%
$\Gamma_{33} K^*(892)^0 \overline{K}_S^0(1430)^0 + \text{c.c.}$	$( 1.5 \pm 0.6 ) \times 10^{-6}$	$\times 10^{-6}$	
$\Gamma_{34} K_1(1270)^\pm K^\mp$	$< 3.22$	$\times 10^{-6}$	CL=90%
$\Gamma_{35} K_1(1400)^\pm K^\mp$	$< 8.3$	$\times 10^{-7}$	CL=90%
$\Gamma_{36} b_1(1235)^\pm \pi^\mp$	$< 4.0$	$\times 10^{-7}$	CL=90%
$\Gamma_{37} \rho \pi$	$< 1.16$	$\times 10^{-6}$	CL=90%
$\Gamma_{38} \pi^+ \pi^- \pi^0$	$< 8.0$	$\times 10^{-7}$	CL=90%
$\Gamma_{39} \omega \pi^0$	$< 1.63$	$\times 10^{-6}$	CL=90%
$\Gamma_{40} \pi^+ \pi^- \pi^0 \pi^0$	$( 1.30 \pm 0.28 ) \times 10^{-5}$	$\times 10^{-5}$	
$\Gamma_{41} K_S^0 K^+ \pi^- + \text{c.c.}$	$( 1.14 \pm 0.33 ) \times 10^{-6}$	$\times 10^{-6}$	
$\Gamma_{42} K^*(892)^0 \overline{K}^0 + \text{c.c.}$	$< 4.22$	$\times 10^{-6}$	CL=90%
$\Gamma_{43} K^*(892)^- K^+ + \text{c.c.}$	$< 1.45$	$\times 10^{-6}$	CL=90%
$\Gamma_{44}$ Sum of 100 exclusive modes	$( 2.90 \pm 0.30 ) \times 10^{-3}$	$\times 10^{-3}$	



See key on page 601

# Meson Particle Listings

## $\tau(2S)$

## Radiative decays

$\Gamma_{45}$	$\gamma\chi_{b1}(1P)$	( 6.9 ± 0.4 ) %		
$\Gamma_{46}$	$\gamma\chi_{b2}(1P)$	( 7.15 ± 0.35 ) %		
$\Gamma_{47}$	$\gamma\chi_{b0}(1P)$	( 3.8 ± 0.4 ) %		
$\Gamma_{48}$	$\gamma f_0(1710)$	< 5.9	$\times 10^{-4}$	CL=90%
$\Gamma_{49}$	$\gamma f'_2(1525)$	< 5.3	$\times 10^{-4}$	CL=90%
$\Gamma_{50}$	$\gamma f_2(1270)$	< 2.41	$\times 10^{-4}$	CL=90%
$\Gamma_{51}$	$\gamma f_J(2220)$			
$\Gamma_{52}$	$\gamma\eta_c(1S)$	< 2.7	$\times 10^{-5}$	CL=90%
$\Gamma_{53}$	$\gamma\chi_{c0}$	< 1.0	$\times 10^{-4}$	CL=90%
$\Gamma_{54}$	$\gamma\chi_{c1}$	< 3.6	$\times 10^{-6}$	CL=90%
$\Gamma_{55}$	$\gamma\chi_{c2}$	< 1.5	$\times 10^{-5}$	CL=90%
$\Gamma_{56}$	$\gamma X(3872) \rightarrow \pi^+\pi^-J/\psi$	< 8	$\times 10^{-7}$	CL=90%
$\Gamma_{57}$	$\gamma X(3872) \rightarrow \pi^+\pi^-\pi^0 J/\psi$	< 2.4	$\times 10^{-6}$	CL=90%
$\Gamma_{58}$	$\gamma X(3915) \rightarrow \omega J/\psi$	< 2.8	$\times 10^{-6}$	CL=90%
$\Gamma_{59}$	$\gamma X(4140) \rightarrow \phi J/\psi$	< 1.2	$\times 10^{-6}$	CL=90%
$\Gamma_{60}$	$\gamma X(4350) \rightarrow \phi J/\psi$	< 1.3	$\times 10^{-6}$	CL=90%
$\Gamma_{61}$	$\gamma\eta_b(1S)$	( 3.9 ± 1.5 ) $\times 10^{-4}$		
$\Gamma_{62}$	$\gamma\eta_b(1S) \rightarrow \gamma$ Sum of 26 exclu- sive modes	< 3.7	$\times 10^{-6}$	CL=90%
$\Gamma_{63}$	$\gamma X_{b\bar{b}} \rightarrow \gamma$ Sum of 26 exclusive modes	< 4.9	$\times 10^{-6}$	CL=90%
$\Gamma_{64}$	$\gamma X \rightarrow \gamma + \geq 4$ prongs	[a] < 1.95	$\times 10^{-4}$	CL=95%
$\Gamma_{65}$	$\gamma A^0 \rightarrow \gamma$ hadrons	< 8	$\times 10^{-5}$	CL=90%
$\Gamma_{66}$	$\gamma a_1^0 \rightarrow \gamma\mu^+\mu^-$	< 8.3	$\times 10^{-6}$	CL=90%

## Lepton Family number (LF) violating modes

$\Gamma_{67}$	$e^\pm\tau^\mp$	LF	< 3.2	$\times 10^{-6}$	CL=90%
$\Gamma_{68}$	$\mu^\pm\tau^\mp$	LF	< 3.3	$\times 10^{-6}$	CL=90%

[a] 1.5 GeV <  $m_X$  < 5.0 GeV

## CONSTRAINED FIT INFORMATION

An overall fit to 3 branching ratios uses 13 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 11.8$  for 11 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle\delta x_i\delta x_j\rangle/(\delta x_i\delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$$x_7 \begin{array}{|c|} \hline x_1 \\ \hline \end{array}$$

 $\tau(2S) \Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$ 

$\Gamma(\mu^+\mu^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_4\Gamma_5/\Gamma$		
VALUE (eV)	DOCUMENT ID	TECN	COMMENT
<b>6.5±1.5±1.0</b>	KOBEL	92	CBAL $e^+e^- \rightarrow \mu^+\mu^-$

 $\Gamma(\tau(1S)\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ 

VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>105.4 ± 1.0 ± 4.2</b>	11.8K	<sup>1</sup> AUBERT	08BP	BABR 10.58 $e^+e^- \rightarrow \gamma\pi^+\pi^-\ell^+\ell^-$
<sup>1</sup> Using B( $\tau(1S) \rightarrow e^+e^-$ ) = (2.38 ± 0.11)% and B( $\tau(1S) \rightarrow \mu^+\mu^-$ ) = (2.48 ± 0.05)%.				

 $\Gamma(\text{hadrons}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ 

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b>0.577 ± 0.009 OUR AVERAGE</b>			
0.581 ± 0.004 ± 0.009	<sup>1</sup> ROSNER	06	CLEO 10.0 $e^+e^- \rightarrow$ hadrons
0.552 ± 0.031 ± 0.017	<sup>1</sup> BARU	96	MD1 $e^+e^- \rightarrow$ hadrons
0.54 ± 0.04 ± 0.02	<sup>1</sup> JAKUBOWSKI	88	CBAL $e^+e^- \rightarrow$ hadrons
0.58 ± 0.03 ± 0.04	<sup>2</sup> GILES	84B	CLEO $e^+e^- \rightarrow$ hadrons
0.60 ± 0.12 ± 0.07	<sup>2</sup> ALBRECHT	82	DASP $e^+e^- \rightarrow$ hadrons
0.54 ± 0.07 $^{+0.09}_{-0.05}$	<sup>2</sup> NICZYPORUK	81C	LENA $e^+e^- \rightarrow$ hadrons
0.41 ± 0.18	<sup>2</sup> BOCK	80	CNTR $e^+e^- \rightarrow$ hadrons

<sup>1</sup> Radiative corrections evaluated following KURAEV 85.<sup>2</sup> Radiative corrections reevaluated by BUCHMUELLER 88 following KURAEV 85. $\tau(2S)$  PARTIAL WIDTHS

$\Gamma(e^+e^-)$	$\Gamma_5$
VALUE (keV)	DOCUMENT ID
<b>0.612 ± 0.011 OUR EVALUATION</b>	

 $\tau(2S)$  BRANCHING RATIOS $\Gamma(\tau(1S)\pi^+\pi^-)/\Gamma_{\text{total}}$ 

Abbreviation MM in the COMMENT field below stands for missing mass.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.85 ± 0.26 OUR FIT</b>				
<b>17.92 ± 0.26 OUR AVERAGE</b>				
16.8 ± 1.1 ± 1.3	906k	<sup>1</sup> LEES	11C	BABR $e^+e^- \rightarrow \pi^+\pi^-\chi$
17.80 ± 0.05 ± 0.37	170k	<sup>2</sup> LEES	11L	BABR $\tau(2S) \rightarrow \pi^+\pi^-\mu^+\mu^-$
18.02 ± 0.02 ± 0.61	851k	<sup>3</sup> BHARI	09	CLEO $e^+e^- \rightarrow \pi^+\pi^-\text{MM}$
17.22 ± 0.17 ± 0.75	11.8K	<sup>4</sup> AUBERT	08BP	BABR $e^+e^- \rightarrow \gamma\pi^+\pi^-\ell^+\ell^-$
19.2 ± 0.2 ± 1.0	52.6k	<sup>5</sup> ALEXANDER	98	CLE2 $\pi^+\pi^-\ell^+\ell^-$ , $\pi^+\pi^-\text{MM}$
18.1 ± 0.5 ± 1.0	11.6k	ALBRECHT	87	ARG $e^+e^- \rightarrow \pi^+\pi^-\text{MM}$
16.9 ± 4.0		GELPHMAN	85	CBAL $e^+e^- \rightarrow e^+e^-\pi^+\pi^-$
19.1 ± 1.2 ± 0.6		BESSION	84	CLEO $\pi^+\pi^-\text{MM}$
18.9 ± 2.6		FONSECA	84	CUSB $e^+e^- \rightarrow \ell^+\ell^-\pi^+\pi^-$
21 ± 7	7	NICZYPORUK	81B	LENA $e^+e^- \rightarrow \ell^+\ell^-\pi^+\pi^-$

<sup>1</sup> LEES 11C reports [ $\Gamma(\tau(2S) \rightarrow \tau(1S)\pi^+\pi^-)/\Gamma_{\text{total}} \times [B(\tau(3S) \rightarrow \tau(2S)\text{anything})] = (1.78 \pm 0.02 \pm 0.11) \times 10^{-2}$  which we divide by our best value B( $\tau(3S) \rightarrow \tau(2S)\text{anything}) = (10.6 \pm 0.8) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>2</sup> Using B( $\tau(1S) \rightarrow \mu^+\mu^-$ ) = (2.48 ± 0.05)%.<sup>3</sup> A weighted average of the inclusive and exclusive results.<sup>4</sup> Using B( $\tau(2S) \rightarrow e^+e^-$ ) = (1.91 ± 0.16)%, B( $\tau(2S) \rightarrow \mu^+\mu^-$ ) = (1.93 ± 0.17)% and,  $\Gamma_{ee}(\tau(2S)) = 0.612 \pm 0.011$  keV.<sup>5</sup> Using B( $\tau(1S) \rightarrow e^+e^-$ ) = (2.52 ± 0.17)% and B( $\tau(1S) \rightarrow \mu^+\mu^-$ ) = (2.48 ± 0.07)%. $\Gamma(\tau(1S)\pi^0\pi^0)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.6 ± 0.4 OUR AVERAGE</b>				
8.43 ± 0.16 ± 0.42	38k	<sup>1</sup> BHARI	09	CLEO $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$
9.2 ± 0.6 ± 0.8	275	<sup>2</sup> ALEXANDER	98	CLE2 $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$
9.5 ± 1.9 ± 1.9	25	ALBRECHT	87	ARG $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$
8.0 ± 1.5		GELPHMAN	85	CBAL $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$
10.3 ± 2.3		FONSECA	84	CUSB $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$

<sup>1</sup> Authors assume B( $\tau(1S) \rightarrow e^+e^-$ ) + B( $\tau(1S) \rightarrow \mu^+\mu^-$ ) = 4.96%.<sup>2</sup> Using B( $\tau(1S) \rightarrow e^+e^-$ ) = (2.52 ± 0.17)% and B( $\tau(1S) \rightarrow \mu^+\mu^-$ ) = (2.48 ± 0.07)%. $\Gamma(\tau(1S)\pi^0\pi^0)/\Gamma(\tau(1S)\pi^+\pi^-)$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.462 ± 0.037	<sup>1</sup> BHARI	09	CLEO $e^+e^- \rightarrow \tau(2S)$

<sup>1</sup> Not independent of other values reported by BHARI 09. $\Gamma(\tau^+\tau^-)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.00 ± 0.21 OUR AVERAGE</b>				
2.00 ± 0.12 ± 0.18	22k	<sup>1</sup> BESSION	07	CLEO $e^+e^- \rightarrow \tau(2S) \rightarrow \tau^+\tau^-$
1.7 ± 1.5 ± 0.6		HAAS	84B	CLEO $e^+e^- \rightarrow \tau^+\tau^-$

<sup>1</sup> BESSION 07 reports [ $\Gamma(\tau(2S) \rightarrow \tau^+\tau^-)/\Gamma_{\text{total}}] / [B(\tau(2S) \rightarrow \mu^+\mu^-)] = 1.04 \pm 0.04 \pm 0.05$  which we multiply by our best value B( $\tau(2S) \rightarrow \mu^+\mu^-$ ) = (1.93 ± 0.17)  $\times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

 $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$ 

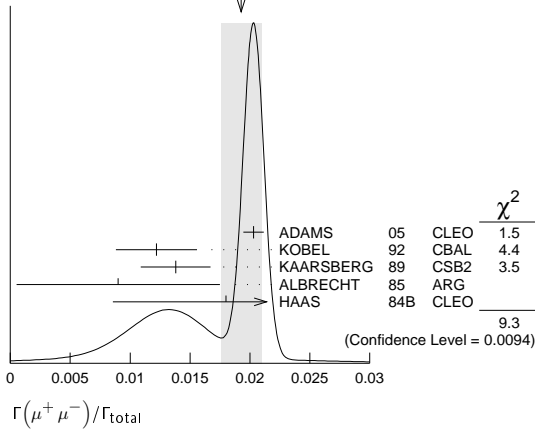
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0193 ± 0.0017 OUR AVERAGE</b>					Error includes scale factor of 2.2. See the ideogram below.
0.0203 ± 0.0003 ± 0.0008		120k	ADAMS	05	CLEO $e^+e^- \rightarrow \mu^+\mu^-$
0.0122 ± 0.0028 ± 0.0019			<sup>1</sup> KOBEL	92	CBAL $e^+e^- \rightarrow \mu^+\mu^-$
0.0138 ± 0.0025 ± 0.0015			KAARSBERG	89	CSB2 $e^+e^- \rightarrow \mu^+\mu^-$
0.009 ± 0.006 ± 0.006			<sup>2</sup> ALBRECHT	85	ARG $e^+e^- \rightarrow \mu^+\mu^-$
0.018 ± 0.008 ± 0.005			HAAS	84B	CLEO $e^+e^- \rightarrow \mu^+\mu^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.038		90	NICZYPORUK	81C	LENA $e^+e^- \rightarrow \mu^+\mu^-$

<sup>1</sup> Taking into account interference between the resonance and continuum.<sup>2</sup> Re-evaluated using B( $\tau(1S) \rightarrow \mu^+\mu^-$ ) = 0.026.

## Meson Particle Listings

 $\Upsilon(2S)$ 

WEIGHTED AVERAGE  
0.0193±0.0017 (Error scaled by 2.2)



$\Gamma(\tau^+\tau^-)/\Gamma(\mu^+\mu^-)$		$\Gamma_3/\Gamma_4$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>1.04±0.04±0.05</b>	22k	BESSION	07 CLEO $e^+e^- \rightarrow \Upsilon(2S)$

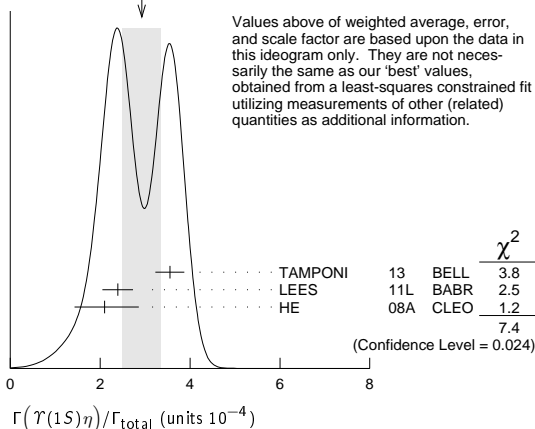
$\Gamma(\Upsilon(1S)\pi^0)/\Gamma_{\text{total}}$		$\Gamma_6/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 4	90	<sup>1</sup> TAMPONI	13 BELL $e^+e^- \rightarrow \Upsilon(1S)\pi^0$
< 18	90	<sup>2</sup> HE	08A CLEO $e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$
< 110	90	ALEXANDER	98 CLE2 $e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$
< 800	90	LURZ	87 CBAL $e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$

<sup>1</sup> TAMPONI 13 reports  $[\Gamma(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^0)/\Gamma_{\text{total}}] / [B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-)] < 2.3 \times 10^{-4}$  which we multiply by our best value  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-) = 17.85 \times 10^{-2}$ .  
<sup>2</sup> Authors assume  $B(\Upsilon(1S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.96\%$ .

$\Gamma(\Upsilon(1S)\pi^0)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$		$\Gamma_6/\Gamma_1$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;2.3</b>	90	TAMPONI	13 BELL $e^+e^- \rightarrow \Upsilon(1S)\pi^0$

$\Gamma(\Upsilon(1S)\eta)/\Gamma_{\text{total}}$		$\Gamma_7/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID TECN COMMENT
<b>2.9 ±0.4 OUR FIT</b>			Error includes scale factor of 2.0.
<b>2.9 ±0.4 OUR AVERAGE</b>			Error includes scale factor of 1.9. See the ideogram below.
2.39±0.31±0.14	112	<sup>1</sup> LEES	11L BABR $\Upsilon(2S) \rightarrow \ell^+\ell^-\eta$
2.1 <sup>+0.7</sup> <sub>-0.6</sub> ±0.3	14	<sup>2</sup> HE	08A CLEO $e^+e^- \rightarrow \ell^+\ell^-\eta$
• • • We use the following data for averages but not for fits. • • •			
3.55±0.32±0.05	241	<sup>3</sup> TAMPONI	13 BELL $e^+e^- \rightarrow \Upsilon(1S)\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 9	90	<sup>1,4</sup> AUBERT	08BP BABR $e^+e^- \rightarrow \gamma\pi^+\pi^-\pi^0\ell^+\ell^-$
< 28	90	ALEXANDER	98 CLE2 $e^+e^- \rightarrow \ell^+\ell^-\eta$
< 50	90	ALBRECHT	87 ARG $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-MM$
< 70	90	LURZ	87 CBAL $e^+e^- \rightarrow \ell^+\ell^-(\gamma\gamma, 3\pi^0)$
< 100	90	BESSION	84 CLEO $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-MM$
< 20	90	FONSECA	84 CUSB $e^+e^- \rightarrow \ell^+\ell^-(\gamma\gamma, \pi^+\pi^-\pi^0)$

WEIGHTED AVERAGE  
2.9±0.4 (Error scaled by 1.9)



<sup>1</sup> Using  $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$  and  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ .

<sup>2</sup> Authors assume  $B(\Upsilon(1S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.96\%$ .

<sup>3</sup> TAMPONI 13 reports  $[\Gamma(\Upsilon(2S) \rightarrow \Upsilon(1S)\eta)/\Gamma_{\text{total}}] / [B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-)] = (1.99 \pm 0.14 \pm 0.11) \times 10^{-3}$  which we multiply by our best value  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-) = (17.85 \pm 0.26) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> Using  $\Gamma_{ee}(\Upsilon(2S)) = 0.612 \pm 0.011$  keV.

### $\Gamma(\Upsilon(1S)\eta)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$ $\Gamma_7/\Gamma_1$

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.64±0.25 OUR FIT</b>					Error includes scale factor of 2.0.
<b>1.99±0.14±0.11</b>		241	TAMPONI	13 BELL	$e^+e^- \rightarrow \Upsilon(1S)\eta$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.35±0.17±0.08		<sup>1</sup> LEES	11L BABR	$\Upsilon(2S) \rightarrow (\pi^+\pi^-)(\gamma\gamma)\mu^+\mu^-$	
< 5.2	90	<sup>2</sup> AUBERT	08BP BABR	$e^+e^- \rightarrow \gamma\pi^+\pi^-(\pi^0)\ell^+\ell^-$	
<sup>1</sup> Not independent of other values reported by LEES 11L.					
<sup>2</sup> Not independent of other values reported by AUBERT 08BP.					

### $\Gamma(\Upsilon(1S)\pi^0)/\Gamma(\Upsilon(1S)\eta)$ $\Gamma_6/\Gamma_7$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.13	90	TAMPONI	13 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^0$

### $\Gamma(\Upsilon/\psi(1S) \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_8/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.006</b>	90	MASCHMANN	90 CBAL	$e^+e^- \rightarrow \text{hadrons}$

### $\Gamma(J/\psi(1S)\eta_c)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.4 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(J/\psi(1S)\chi_{c0})/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.4 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(J/\psi(1S)\chi_{c1})/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.2 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(J/\psi(1S)\chi_{c2})/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.0 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(J/\psi(1S)\eta_c(2S))/\Gamma_{\text{total}}$ $\Gamma_{13}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(J/\psi(1S)X(3940))/\Gamma_{\text{total}}$ $\Gamma_{14}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.0 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(J/\psi(1S)X(4160))/\Gamma_{\text{total}}$ $\Gamma_{15}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.0 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow J/\psi X$

### $\Gamma(\psi(2S)\eta_c)/\Gamma_{\text{total}}$ $\Gamma_{16}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.1 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

### $\Gamma(\psi(2S)\chi_{c0})/\Gamma_{\text{total}}$ $\Gamma_{17}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.7 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

### $\Gamma(\psi(2S)\chi_{c1})/\Gamma_{\text{total}}$ $\Gamma_{18}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

### $\Gamma(\psi(2S)\chi_{c2})/\Gamma_{\text{total}}$ $\Gamma_{19}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.9 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

### $\Gamma(\psi(2S)\eta_c(2S))/\Gamma_{\text{total}}$ $\Gamma_{20}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.3 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

### $\Gamma(\psi(2S)X(3940))/\Gamma_{\text{total}}$ $\Gamma_{21}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.9 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

### $\Gamma(\psi(2S)X(4160))/\Gamma_{\text{total}}$ $\Gamma_{22}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.9 × 10<sup>-6</sup></b>	90	YANG	14 BELL	$e^+e^- \rightarrow \psi(2S) X$

See key on page 601

Meson Particle Listings  
 $\Upsilon(2S)$ 

$\Gamma(\overline{2H} \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{23}/\Gamma$
VALUE (units $10^{-5}$ )	EVTS

**$2.78^{+0.30}_{-0.26}$  OUR AVERAGE** Error includes scale factor of 1.2.

$2.64 \pm 0.11^{+0.26}_{-0.21}$	LEES	14G	BABR	$e^+e^- \rightarrow \overline{2H} X$
$3.37 \pm 0.50 \pm 0.25$	58	ASNER	07	CLEO $e^+e^- \rightarrow \overline{2H} X$

$\Gamma(ggg)/\Gamma_{\text{total}}$				$\Gamma_{25}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>58.8 \pm 1.2</math></b>	6M	<sup>1</sup> BESSON	06A	CLEO
				$\Upsilon(2S) \rightarrow \text{hadrons}$

<sup>1</sup> Calculated using the value  $\Gamma(\gamma gg)/\Gamma(ggg) = (3.18 \pm 0.04 \pm 0.22 \pm 0.41)\%$  from BESSON 06A and PDG 08 values of  $B(\pi^+\pi^-\Upsilon(1S)) = (18.1 \pm 0.4)\%$ ,  $B(\pi^0\pi^0\Upsilon(1S)) = (8.6 \pm 0.4)\%$ ,  $B(\mu^+\mu^-\Upsilon(1S)) = (1.93 \pm 0.17)\%$ , and  $R_{\text{hadrons}} = 3.51$ . The statistical error is negligible and the systematic error is partially correlated with that of  $\Gamma(\gamma gg)/\Gamma_{\text{total}}$  measurement of BESSON 06A.

$\Gamma(\gamma gg)/\Gamma(ggg)$					$\Gamma_{26}/\Gamma_{25}$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>3.18 \pm 0.04 \pm 0.47</math></b>	6M	BESSON	06A	CLEO	$\Upsilon(2S) \rightarrow (\gamma +) \text{ hadrons}$

$\Gamma(\phi K^+ K^-)/\Gamma_{\text{total}}$				$\Gamma_{27}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.58 \pm 0.33 \pm 0.18</math></b>	58	SHEN	12A	BELL
				$\Upsilon(1S) \rightarrow 2(K^+ K^-)$

$\Gamma(\omega\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;2.58</b>	90	SHEN	12A	BELL	$\Upsilon(1S) \rightarrow 2(\pi^+\pi^-)\pi^0$

$\Gamma(K^*(892)^0 K^- \pi^+ + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>2.32 \pm 0.40 \pm 0.54</math></b>	135	SHEN	12A	BELL	$\Upsilon(1S) \rightarrow K^+ K^- \pi^+ \pi^-$

$\Gamma(\phi f_2'(1525))/\Gamma_{\text{total}}$				$\Gamma_{30}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.33</b>	90	SHEN	12A	BELL
				$\Upsilon(1S) \rightarrow 2(K^+ K^-)$

$\Gamma(\omega f_2(1270))/\Gamma_{\text{total}}$				$\Gamma_{31}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.57</b>	90	SHEN	12A	BELL
				$\Upsilon(1S) \rightarrow 2(\pi^+ \pi^-) \pi^0$

$\Gamma(\rho(770)a_2(1320))/\Gamma_{\text{total}}$				$\Gamma_{32}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.88</b>	90	SHEN	12A	BELL
				$\Upsilon(1S) \rightarrow 2(\pi^+ \pi^-) \pi^0$

$\Gamma(K^*(892)^0 \overline{K}_2^*(1430)^0 + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.53 \pm 0.52 \pm 0.19</math></b>	32	SHEN	12A	BELL	$\Upsilon(1S) \rightarrow K^+ K^- \pi^+ \pi^-$

$\Gamma(K_1(1270)^\pm K^\mp)/\Gamma_{\text{total}}$				$\Gamma_{34}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.22</b>	90	SHEN	12A BELL	$\Upsilon(1S) \rightarrow K^+ K^- \pi^+ \pi^-$

$\Gamma(K_1(1400)^\pm K^\mp)/\Gamma_{\text{total}}$				$\Gamma_{35}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.83</b>	90	SHEN	12A	BELL
				$\Upsilon(1S) \rightarrow K^+ K^- \pi^+ \pi^-$

$\Gamma(b_1(1235)^\pm \pi^\mp)/\Gamma_{\text{total}}$				$\Gamma_{36}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.40</b>	90	SHEN	12A	BELL
				$\Upsilon(1S) \rightarrow 2(\pi^+ \pi^-) \pi^0$

$\Gamma(\rho\pi)/\Gamma_{\text{total}}$				$\Gamma_{37}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.16</b>	90	SHEN	13	BELL
				$\Upsilon(2S) \rightarrow \pi^+ \pi^- \pi^0$

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{38}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.80</b>	90	SHEN	13	BELL	$\Upsilon(2S) \rightarrow \pi^+\pi^-\pi^0$

$\Gamma(\omega\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{39}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.63</b>	90	SHEN	13	BELL
				$\Upsilon(2S) \rightarrow \pi^+ \pi^- \pi^0 \pi^0$

$\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{40}/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>13.0 \pm 1.9 \pm 2.1</math></b>	$261 \pm 37$	SHEN	13	BELL
				$\Upsilon(2S) \rightarrow \pi^+\pi^-\pi^0\pi^0$

$\Gamma(K_S^0 K^+ \pi^- + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{41}/\Gamma$
VALUE (units $10^{-6}$ )	CL%
<b><math>1.14 \pm 0.30 \pm 0.13</math></b>	$40 \pm 10$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 $<3.2$  90 <sup>1</sup> DOBBS 12A  $\Upsilon(2S) \rightarrow K_S^0 K^- \pi^+$   
<sup>1</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.

$\Gamma(K^*(892)^0 \bar{K}^0 + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_{42}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.22</b>	90	SHEN	13	BELL
				$\Upsilon(2S) \rightarrow K_S^0 K^- \pi^+$

$\Gamma(K^*(892)^- K^+ + \text{c.c.})/\Gamma_{\text{total}}$				$\Gamma_{43}/\Gamma$
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.45</b>	90	SHEN	13	BELL
				$\Upsilon(2S) \rightarrow K_S^0 K^- \pi^+$

$\Gamma(\text{Sum of 100 exclusive modes})/\Gamma_{\text{total}}$	$\Gamma_{44}/\Gamma$
VALUE (units $10^{-2}$ )	EVTS
<b><math>0.29 \pm 0.03</math></b>	$1,2$ DOBBS 12A $\Upsilon(2S) \rightarrow \text{hadrons}$

<sup>1</sup> DOBBS 12A presents individual exclusive branching fractions or upper limits for 100 modes of four to ten pions, kaons, or protons.  
<sup>2</sup> Obtained by analyzing CLEO III data but not authored by the CLEO Collaboration.

$\Gamma(\gamma\chi_{b1}(1P))/\Gamma_{\text{total}}$				$\Gamma_{45}/\Gamma$	
VALUE		EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.069 \pm 0.004</math></b>	<b>OUR AVERAGE</b>				
$0.0693 \pm 0.0012 \pm 0.0041$	407k		ARTUSO	05	CLEO $e^+e^- \rightarrow \gamma X$
$0.069 \pm 0.005 \pm 0.009$			EDWARDS	99	CLE2 $\Upsilon(2S) \rightarrow \gamma\chi(1P)$
$0.091 \pm 0.018 \pm 0.022$			ALBRECHT	85E	ARG $e^+e^- \rightarrow \gamma \text{conv. } X$
$0.065 \pm 0.007 \pm 0.012$			NERNST	85	CBAL $e^+e^- \rightarrow \gamma X$
$0.080 \pm 0.017 \pm 0.016$			HAAS	84	CLEO $e^+e^- \rightarrow \gamma \text{conv. } X$
$0.059 \pm 0.014$			KLOPFEN...	83	CUSB $e^+e^- \rightarrow \gamma X$

$\Gamma(\gamma\chi_{b2}(1P))/\Gamma_{\text{total}}$				$\Gamma_{46}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.0715 \pm 0.0035</math> OUR AVERAGE</b>				
$0.0724 \pm 0.0011 \pm 0.0040$	410k	ARTUSO	05	CLEO $e^+e^- \rightarrow \gamma X$
$0.074 \pm 0.005 \pm 0.008$		EDWARDS	99	CLE2 $\Upsilon(2S) \rightarrow \gamma\chi(1P)$
$0.098 \pm 0.021 \pm 0.024$		ALBRECHT	85E	ARG $e^+e^- \rightarrow \gamma \text{conv. } X$
$0.058 \pm 0.007 \pm 0.010$		NERNST	85	CBAL $e^+e^- \rightarrow \gamma X$
$0.102 \pm 0.018 \pm 0.021$		HAAS	84	CLEO $e^+e^- \rightarrow \gamma \text{conv. } X$
$0.061 \pm 0.014$		KLOPFEN...	83	CUSB $e^+e^- \rightarrow \gamma X$

$\Gamma(\gamma\chi_{b0}(1P))/\Gamma_{\text{total}}$				$\Gamma_{47}/\Gamma$	
VALUE		EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.038 \pm 0.004</math></b>	<b>OUR AVERAGE</b>				
$0.0375 \pm 0.0012 \pm 0.0047$	198k		ARTUSO	05	CLEO $e^+e^- \rightarrow \gamma X$
$0.034 \pm 0.005 \pm 0.006$			EDWARDS	99	CLE2 $\Upsilon(2S) \rightarrow \gamma\chi(1P)$
$0.064 \pm 0.014 \pm 0.016$			ALBRECHT	85E	ARG $e^+e^- \rightarrow \gamma\text{conv. } X$
$0.036 \pm 0.008 \pm 0.009$			NERNST	85	CBAL $e^+e^- \rightarrow \gamma X$
$0.044 \pm 0.023 \pm 0.009$			HAAS	84	CLEO $e^+e^- \rightarrow \gamma\text{conv. } X$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$0.035 \pm 0.014$			KLOPFEN...	83	CUSB $e^+e^- \rightarrow \gamma X$

$\Gamma(\gamma f_0(1710))/\Gamma_{\text{total}}$					$\Gamma_{48}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;59</b>	90	<sup>1</sup> ALBRECHT 89	ARG	$\Upsilon(2S) \rightarrow \gamma K^+ K^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 5.9	90	<sup>2</sup> ALBRECHT 89	ARG	$\Upsilon(2S) \rightarrow \gamma \pi^+ \pi^-$	
<sup>1</sup> Re-evaluated assuming $B(f_0(1710) \rightarrow K^+ K^-) = 0.19$ .					
<sup>2</sup> Includes unknown branching ratio of $f_0(1710) \rightarrow \pi^+ \pi^-$ .					

$\Gamma(\gamma f_2'(1525))/\Gamma_{\text{total}}$				$\Gamma_{49}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;53</b>	90	<sup>1</sup> ALBRECHT 89	ARG	$\Upsilon(2S) \rightarrow \gamma K^+ K^-$
<sup>1</sup> Re-evaluated assuming $B(f_2'(1525) \rightarrow K \overline{K}) = 0.71$ .				

$\Gamma(\gamma f_2(1270))/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;24.1</b>	90	<sup>1</sup> ALBRECHT 89	ARG	$T(2S) \rightarrow \gamma \pi^+ \pi^-$	
<sup>1</sup> Using $B(f_2(1270) \rightarrow \pi\pi) = 0.84$ .					

$\Gamma(\gamma f_J(2220))/\Gamma_{\text{total}}$				$\Gamma_{51}/\Gamma$
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<6.8$	90	<sup>1</sup> ALBRECHT 89 ARG	$\Upsilon(2S) \rightarrow \gamma K^+ K^-$	
<sup>1</sup> Includes unknown branching ratio of $f_J(2220) \rightarrow K^+ K^-$ .				

$\Gamma(\gamma\eta_c(1S))/\Gamma_{\text{total}}$			$\Gamma_{52}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN COMMENT
$<2.7 \times 10^{-5}$	90	WANG 11B	BELL $\Upsilon(2S) \rightarrow \gamma X$



See key on page 601

## Meson Particle Listings

 $\Upsilon(1D)$ ,  $\chi_{b0}(2P)$  $\Upsilon(1D)$  BRANCHING RATIOS

$\Gamma(\eta \Upsilon(1S))/\Gamma(\gamma\gamma \Upsilon(1S))$		$\Gamma_3/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.25</b>	90	BONVICINI	04 CLE3 $\Upsilon(3S) \rightarrow 4\gamma \ell^+ \ell^-$

$\Gamma(\pi^+ \pi^- \Upsilon(1S))/\Gamma_{\text{total}}$		$\Gamma_4/\Gamma$	
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b><math>0.66^{+0.15}_{-0.14} \pm 0.06</math></b>		<sup>1</sup> DEL-AMO-SA..10R	BABR $\Upsilon(3S) \rightarrow \gamma\gamma \pi^+ \pi^- \ell^+ \ell^-$

<sup>1</sup> Using theoretical predictions for  $B(\chi_{bJ}(2P) \rightarrow \gamma \Upsilon(1D))$ .

$\Gamma(\pi^+ \pi^- \Upsilon(1S))/\Gamma(\gamma\gamma \Upsilon(1S))$		$\Gamma_4/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;1.2</b>	90	<sup>2</sup> BONVICINI	04 CLE3 $\Upsilon(3S) \rightarrow 4\gamma \ell^+ \ell^-$

<sup>2</sup> Assuming  $J = 2$ . $\Upsilon(1D)$  REFERENCES

DEL-AMO-SA..10R	PR D82 111102	P. del Amo Sanchez et al.	(BABAR Collab.)
BONVICINI 04	PR D70 032001	G. Bonvicini et al.	(CLEO Collab.)

$$\chi_{b0}(2P) \quad J^G(J^{PC}) = 0^+(0^{++})$$

$J$  needs confirmation.

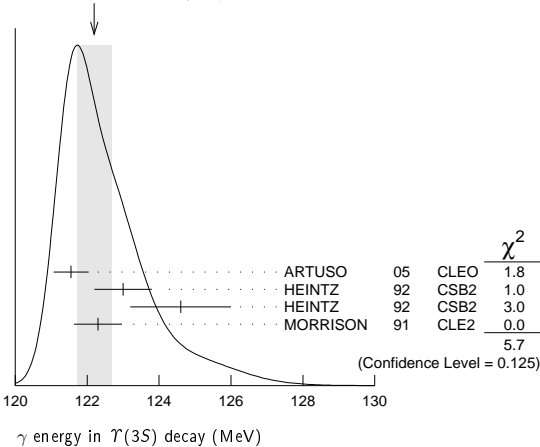
Observed in radiative decay of the  $\Upsilon(3S)$ , therefore  $C = +$ . Branching ratio requires E1 transition, M1 is strongly disfavored, therefore  $P = +$ .

 $\chi_{b0}(2P)$  MASS

VALUE (MeV)	DOCUMENT ID
<b><math>10232.5 \pm 0.4 \pm 0.5</math> OUR EVALUATION</b>	From $\gamma$ energy below, using $\Upsilon(3S)$ mass = $10355.2 \pm 0.5$ MeV

 $\gamma$  ENERGY IN  $\Upsilon(3S)$  DECAY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN COMMENT
<b><math>121.9 \pm 0.4</math> OUR EVALUATION</b>		Treating systematic errors as correlated	
<b><math>122.2 \pm 0.5</math> OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.	
$121.55 \pm 0.16 \pm 0.46$		ARTUSO	05 CLEO $\Upsilon(3S) \rightarrow \gamma X$
$123.0 \pm 0.8$	4959	<sup>1</sup> HEINTZ	92 CSB2 $e^+ e^- \rightarrow \gamma X$
$124.6 \pm 1.4$	17	<sup>2</sup> HEINTZ	92 CSB2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma\gamma$
$122.3 \pm 0.3 \pm 0.6$	9903	MORRISON	91 CLE2 $e^+ e^- \rightarrow \gamma X$

<sup>1</sup> A systematic uncertainty on the energy scale of 0.9% not included. Supersedes NARAIN 91.<sup>2</sup> A systematic uncertainty on the energy scale of 0.9% not included. Supersedes HEINTZ 91.WEIGHTED AVERAGE  
 $122.2 \pm 0.5$  (Error scaled by 1.4) $\chi_{b0}(2P)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \quad \gamma \Upsilon(2S)$	$(4.6 \pm 2.1) \%$	
$\Gamma_2 \quad \gamma \Upsilon(1S)$	$(9 \pm 6) \times 10^{-3}$	
$\Gamma_3 \quad D^0 X$	$< 8.2 \%$	90%
$\Gamma_4 \quad \pi^+ \pi^- K^+ K^- \pi^0$	$< 3.4 \times 10^{-5}$	90%
$\Gamma_5 \quad 2\pi^+ \pi^- K^- K_S^0$	$< 5 \times 10^{-5}$	90%
$\Gamma_6 \quad 2\pi^+ \pi^- K^- K_S^0 2\pi^0$	$< 2.2 \times 10^{-4}$	90%
$\Gamma_7 \quad 2\pi^+ 2\pi^- 2\pi^0$	$< 2.4 \times 10^{-4}$	90%

$\Gamma_8 \quad 2\pi^+ 2\pi^- K^+ K^-$	$< 1.5 \times 10^{-4}$	90%
$\Gamma_9 \quad 2\pi^+ 2\pi^- K^+ K^- \pi^0$	$< 2.2 \times 10^{-4}$	90%
$\Gamma_{10} \quad 2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	$< 1.1 \times 10^{-3}$	90%
$\Gamma_{11} \quad 3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	$< 7 \times 10^{-4}$	90%
$\Gamma_{12} \quad 3\pi^+ 3\pi^-$	$< 7 \times 10^{-5}$	90%
$\Gamma_{13} \quad 3\pi^+ 3\pi^- 2\pi^0$	$< 1.2 \times 10^{-3}$	90%
$\Gamma_{14} \quad 3\pi^+ 3\pi^- K^+ K^-$	$< 1.5 \times 10^{-4}$	90%
$\Gamma_{15} \quad 3\pi^+ 3\pi^- K^+ K^- \pi^0$	$< 7 \times 10^{-4}$	90%
$\Gamma_{16} \quad 4\pi^+ 4\pi^-$	$< 1.7 \times 10^{-4}$	90%
$\Gamma_{17} \quad 4\pi^+ 4\pi^- 2\pi^0$	$< 6 \times 10^{-4}$	90%

 $\chi_{b0}(2P)$  BRANCHING RATIOS

$\Gamma(\gamma \Upsilon(2S))/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b><math>0.046 \pm 0.020 \pm 0.007</math></b>		<sup>3</sup> HEINTZ	92 CSB2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.028$	90	<sup>4</sup> LEES	11J BABR $\Upsilon(3S) \rightarrow X\gamma$
$<0.089$	90	<sup>5</sup> CRAWFORD	92B CLE2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma\gamma$

<sup>3</sup> Using  $B(\Upsilon(2S) \rightarrow \mu^+ \mu^-) = (1.44 \pm 0.10)\%$ ,  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = (6.0 \pm 0.4 \pm 0.6)\%$  and assuming  $e\mu$  universality. Supersedes HEINTZ 91.<sup>4</sup> LEES 11J quotes a central value of  $\Gamma(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(2S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))/\Gamma_{\text{total}} = (-0.3 \pm 0.2^{+0.5}_{-0.4})\%$ .<sup>5</sup> Using  $B(\Upsilon(2S) \rightarrow \mu^+ \mu^-) = (1.37 \pm 0.26)\%$ ,  $B(\Upsilon(3S) \rightarrow \gamma\gamma \Upsilon(2S)) \times 2 B(\Upsilon(2S) \rightarrow \mu^+ \mu^-) < 1.19 \times 10^{-4}$ , and  $B(\Upsilon(3S) \rightarrow \chi_{b0}(2P)\gamma) = 0.049$ .

$\Gamma(\gamma \Upsilon(1S))/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b><math>0.009 \pm 0.006 \pm 0.001</math></b>		<sup>6</sup> HEINTZ	92 CSB2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.012$	90	<sup>7</sup> LEES	11J BABR $\Upsilon(3S) \rightarrow X\gamma$
$<0.025$	90	<sup>8</sup> CRAWFORD	92B CLE2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma\gamma$

<sup>6</sup> Using  $B(\Upsilon(1S) \rightarrow \mu^+ \mu^-) = (2.57 \pm 0.07)\%$ ,  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = (6.0 \pm 0.4 \pm 0.6)\%$  and assuming  $e\mu$  universality. Supersedes HEINTZ 91.<sup>7</sup> LEES 11J quotes a central value of  $\Gamma(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))/\Gamma_{\text{total}} = (3.9 \pm 2.2^{+1.2}_{-0.6}) \times 10^{-4}$ .<sup>8</sup> Using  $B(\Upsilon(1S) \rightarrow \mu^+ \mu^-) = (2.57 \pm 0.07)\%$ ,  $B(\Upsilon(3S) \rightarrow \gamma\gamma \Upsilon(1S)) \times 2 B(\Upsilon(1S) \rightarrow \mu^+ \mu^-) < 0.63 \times 10^{-4}$ , and  $B(\Upsilon(3S) \rightarrow \chi_{b0}(2P)\gamma) = 0.049$ .

$\Gamma(D^0 X)/\Gamma_{\text{total}}$		$\Gamma_3/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b><math>&lt;8.2 \times 10^{-2}</math></b>	90	<sup>9,10</sup> BRIERE	08 CLEO $\Upsilon(3S) \rightarrow \gamma D^0 X$

<sup>9</sup> For  $p_{D^0} > 2.5$  GeV/c.<sup>10</sup> The authors also present their result as  $(4.1 \pm 3.0 \pm 0.4) \times 10^{-2}$ .

$\Gamma(\pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$		$\Gamma_4/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.34</b>	90	<sup>11</sup> ASNER	08A CLEO $\Upsilon(3S) \rightarrow \gamma \pi^+ \pi^- K^+ K^- \pi^0$

<sup>11</sup> ASNER 08A reports  $[\Gamma(\chi_{b0}(2P) \rightarrow \pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))]$   $< 2 \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .

$\Gamma(2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}$		$\Gamma_5/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.5</b>	90	<sup>12</sup> ASNER	08A CLEO $\Upsilon(3S) \rightarrow \gamma 2\pi^+ \pi^- K^- K_S^0$

<sup>12</sup> ASNER 08A reports  $[\Gamma(\chi_{b0}(2P) \rightarrow 2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))]$   $< 3 \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .

$\Gamma(2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}$		$\Gamma_6/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;2.2</b>	90	<sup>13</sup> ASNER	08A CLEO $\Upsilon(3S) \rightarrow \gamma 2\pi^+ \pi^- K^- 2\pi^0$

<sup>13</sup> ASNER 08A reports  $[\Gamma(\chi_{b0}(2P) \rightarrow 2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))]$   $< 13 \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .

$\Gamma(2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}$		$\Gamma_7/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;2.4</b>	90	<sup>14</sup> ASNER	08A CLEO $\Upsilon(3S) \rightarrow \gamma 2\pi^+ 2\pi^- 2\pi^0$

<sup>14</sup> ASNER 08A reports  $[\Gamma(\chi_{b0}(2P) \rightarrow 2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))]$   $< 14 \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .

$\Gamma(2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}$		$\Gamma_8/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;1.5</b>	90	<sup>15</sup> ASNER	08A CLEO $\Upsilon(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^-$

<sup>15</sup> ASNER 08A reports  $[\Gamma(\chi_{b0}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P))]$   $< 9 \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(3S) \rightarrow \gamma \chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .

# Meson Particle Listings

## $\chi_{b0}(2P)$ , $\chi_{b1}(2P)$

### $\Gamma(2\pi^+2\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.2	90	16 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 2\pi^+2\pi^-K^+K^-\pi^0$
16 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 2\pi^+2\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $13 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(2\pi^+2\pi^-K^+K^-2\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<11	90	17 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 2\pi^+2\pi^-K^+K^-2\pi^0$
17 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 2\pi^+2\pi^-K^+K^-2\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $63 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(3\pi^+2\pi^-K^-K_S^0\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<7	90	18 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+2\pi^-K^-K_S^0\pi^0$
18 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 3\pi^+2\pi^-K^-K_S^0\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $39 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<0.7	90	19 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+3\pi^-$
19 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 3\pi^+3\pi^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $4 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-2\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{13}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<12	90	20 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+3\pi^-2\pi^0$
20 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 3\pi^+3\pi^-2\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $72 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-K^+K^-)/\Gamma_{\text{total}}$ $\Gamma_{14}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.5	90	21 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+3\pi^-K^+K^-$
21 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 3\pi^+3\pi^-K^+K^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $9 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{15}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<7	90	22 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+3\pi^-K^+K^-\pi^0$
22 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $43 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(4\pi^+4\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{16}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.7	90	23 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 4\pi^+4\pi^-$
23 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 4\pi^+4\pi^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $10 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(4\pi^+4\pi^-2\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{17}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<6	90	24 ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 4\pi^+4\pi^-2\pi^0$
24 ASNER 08A reports $[\Gamma(\chi_{b0}(2P) \rightarrow 4\pi^+4\pi^-2\pi^0)/\Gamma_{\text{total}}] \times [B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))]$ < $38 \times 10^{-6}$ which we divide by our best value $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = 5.9 \times 10^{-2}$ .				

### $\Gamma(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))/\Gamma_{\text{total}}$ $\Gamma_2/\Gamma \times \Gamma_{22}^{\Upsilon(3S)}/\Gamma \Upsilon(3S)$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<8.2	90	25 LEES	11J BABR	$\Upsilon(3S) \rightarrow X\gamma$
25 LEES 11J quotes a central value of $\Gamma(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))/\Gamma_{\text{total}} = (3.9 \pm 2.2^{+1.2}_{-0.6}) \times 10^{-4}$ and derives a 90% CL upper limit of $B(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(1S)) < 1.2\%$ using $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = (5.9 \pm 0.6)\%$ .				

### $\Gamma(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(2S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))/\Gamma_{\text{total}}$ $\Gamma_1/\Gamma \times \Gamma_{22}^{\Upsilon(3S)}/\Gamma \Upsilon(3S)$

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<1.6	90	26 LEES	11J BABR	$\Upsilon(3S) \rightarrow X\gamma$
26 LEES 11J quotes a central value of $\Gamma(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(2S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P))/\Gamma_{\text{total}} = (-0.3 \pm 0.2^{+0.5}_{-0.4})\%$ and derives a 90% CL upper limit of $B(\chi_{b0}(2P) \rightarrow \gamma \Upsilon(2S)) < 2.8\%$ using $B(\Upsilon(3S) \rightarrow \gamma\chi_{b0}(2P)) = (5.9 \pm 0.6)\%$ .				

## $\chi_{b0}(2P)$ REFERENCES

LEES	11J	PR D84 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ASNER	08A	PR D78 091103	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	08	PR D78 092007	R.A. Briere <i>et al.</i>	(CLEO Collab.)
ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
CRAWFORD	92B	PL B294 139	G. Crawford, R. Fulton	(CLEO Collab.)
HEINTZ	92	PR D46 1928	U. Heintz <i>et al.</i>	(CUSB II Collab.)
HEINTZ	91	PRL 66 1563	U. Heintz <i>et al.</i>	(CUSB Collab.)
MORRISON	91	PRL 67 1696	R.J. Morrison <i>et al.</i>	(CLEO Collab.)
NARAIN	91	PRL 66 3113	M. Narain <i>et al.</i>	(CUSB Collab.)

## $\chi_{b1}(2P)$

$$J^G(J^{PC}) = 0^+(1^++)$$

$J$  needs confirmation.

Observed in radiative decay of the  $\Upsilon(3S)$ , therefore  $C = +$ . Branching ratio requires E1 transition, M1 is strongly disfavored, therefore  $P = +$ .

## $\chi_{b1}(2P)$ MASS

VALUE (MeV)	DOCUMENT ID
<b><math>10255.46 \pm 0.22 \pm 0.50</math> OUR EVALUATION</b>	From $\gamma$ energy below, using $\Upsilon(3S)$ mass = $10355.2 \pm 0.5$ MeV

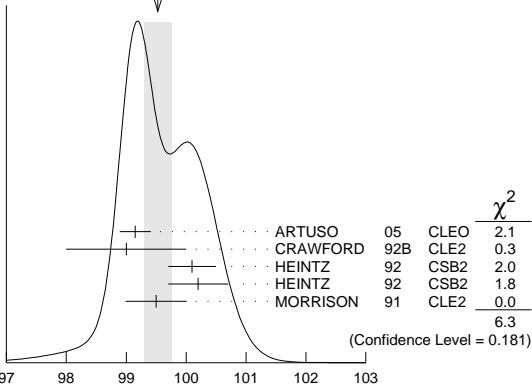
## $m_{\chi_{b1}(2P)} - m_{\chi_{b0}(2P)}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>23.5 \pm 0.7 \pm 0.7</math></b>	1 HEINTZ	92 CSB2	$e^+e^- \rightarrow \gamma X, \ell^+\ell^-\gamma\gamma$
1 From the average photon energy for inclusive and exclusive events. Supersedes NARAIN 91.			

## $\gamma$ ENERGY IN $\Upsilon(3S)$ DECAY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>99.26 \pm 0.22</math> OUR EVALUATION</b>		Treating systematic errors as correlated		
<b><math>99.53 \pm 0.23</math> OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
$99.15 \pm 0.07 \pm 0.25$		ARTUSO 05	CLEO	$\Upsilon(3S) \rightarrow \gamma X$
$99 \pm 1$	169	CRAWFORD 92B	CLE2	$e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$
$100.1 \pm 0.4$	11147	2 HEINTZ 92	CSB2	$e^+e^- \rightarrow \gamma X$
$100.2 \pm 0.5$	223	3 HEINTZ 92	CSB2	$e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$
$99.5 \pm 0.1 \pm 0.5$	25759	MORRISON 91	CLE2	$e^+e^- \rightarrow \gamma X$
2 A systematic uncertainty on the energy scale of 0.9% not included. Supersedes NARAIN 91.				
3 A systematic uncertainty on the energy scale of 0.9% not included. Supersedes HEINTZ 91.				

WEIGHTED AVERAGE  
 $99.53 \pm 0.23$  (Error scaled by 1.3)



## $\chi_{b1}(2P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor
$\Gamma_1 \quad \omega \Upsilon(1S)$	$(1.63^{+0.40}_{-0.34})\%$	
$\Gamma_2 \quad \gamma \Upsilon(2S)$	$(19.9 \pm 1.9)\%$	
$\Gamma_3 \quad \gamma \Upsilon(1S)$	$(9.2 \pm 0.8)\%$	
$\Gamma_4 \quad \pi\pi\chi_{b1}(1P)$	$(9.1 \pm 1.3) \times 10^{-3}$	1.1
$\Gamma_5 \quad D^0 X$	$(8.8 \pm 1.7)\%$	
$\Gamma_6 \quad \pi^+\pi^-K^+K^-\pi^0$	$(3.1 \pm 1.0) \times 10^{-4}$	
$\Gamma_7 \quad 2\pi^+\pi^-K^-K_S^0$	$(1.1 \pm 0.5) \times 10^{-4}$	
$\Gamma_8 \quad 2\pi^+\pi^-K^-K_S^0\pi^0$	$(7.7 \pm 3.2) \times 10^{-4}$	
$\Gamma_9 \quad 2\pi^+2\pi^-2\pi^0$	$(5.9 \pm 2.0) \times 10^{-4}$	

$\Gamma_{10}$	$2\pi^+ 2\pi^- K^+ K^-$	$(10 \pm 4) \times 10^{-5}$
$\Gamma_{11}$	$2\pi^+ 2\pi^- K^+ K^- \pi^0$	$(5.5 \pm 1.8) \times 10^{-4}$
$\Gamma_{12}$	$2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	$(10 \pm 4) \times 10^{-4}$
$\Gamma_{13}$	$3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	$(6.7 \pm 2.6) \times 10^{-4}$
$\Gamma_{14}$	$3\pi^+ 3\pi^-$	$(1.2 \pm 0.4) \times 10^{-4}$
$\Gamma_{15}$	$3\pi^+ 3\pi^- 2\pi^0$	$(1.2 \pm 0.4) \times 10^{-3}$
$\Gamma_{16}$	$3\pi^+ 3\pi^- K^+ K^-$	$(2.0 \pm 0.8) \times 10^{-4}$
$\Gamma_{17}$	$3\pi^+ 3\pi^- K^+ K^- \pi^0$	$(6.1 \pm 2.2) \times 10^{-4}$
$\Gamma_{18}$	$4\pi^+ 4\pi^-$	$(1.7 \pm 0.6) \times 10^{-4}$
$\Gamma_{19}$	$4\pi^+ 4\pi^- 2\pi^0$	$(1.9 \pm 0.7) \times 10^{-3}$

 $\chi_{b1}(2P)$  BRANCHING RATIOS

$\Gamma(\omega \mathcal{T}(1S))/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.63 \pm 0.35 \pm 0.16</math> <math>-0.31 -0.15</math></b>	$32.6^{+6.9}_{-6.1}$	4	CRONIN-HEN..04	CLE3 $\mathcal{T}(3S) \rightarrow \gamma \omega \mathcal{T}(1S)$
<sup>4</sup> Using $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (11.3 \pm 0.6)\%$ and $B(\mathcal{T}(1S) \rightarrow \ell^+ \ell^-) = 2 B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = 2 (2.48 \pm 0.06)\%$ .				

$\Gamma(\gamma \mathcal{T}(2S))/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.199 \pm 0.019</math> OUR AVERAGE</b>				
$0.190 \pm 0.018 \pm 0.017$	4.3k	5	LEES	11J BABR $\mathcal{T}(3S) \rightarrow X \gamma$
$0.356 \pm 0.042 \pm 0.092$		6	CRAWFORD	92B CLE2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma \gamma$
$0.199 \pm 0.020 \pm 0.022$		7	HEINTZ	92 CSB2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma \gamma$
<sup>5</sup> LEES 11J reports $[\Gamma(\chi_{b1}(2P) \rightarrow \gamma \mathcal{T}(2S))/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(2.4 \pm 0.1 \pm 0.2) \times 10^{-2}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>6</sup> Using $B(\mathcal{T}(2S) \rightarrow \mu^+ \mu^-) = (1.37 \pm 0.26)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \gamma \mathcal{T}(2S)) \times 2 B(\mathcal{T}(2S) \rightarrow \mu^+ \mu^-) = (10.23 \pm 1.20 \pm 1.26) \times 10^{-4}$ , and $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = 0.105^{+0.003}_{-0.002} \pm 0.013$ .				
<sup>7</sup> Using $B(\mathcal{T}(2S) \rightarrow \mu^+ \mu^-) = (1.44 \pm 0.10)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (11.5 \pm 0.5 \pm 0.5)\%$ and assuming $e\mu$ universality. Supersedes HEINTZ 91.				

$\Gamma(\gamma \mathcal{T}(1S))/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.092 \pm 0.008</math> OUR AVERAGE</b>				
$0.098 \pm 0.005 \pm 0.009$	15k	8	LEES	11J BABR $\mathcal{T}(3S) \rightarrow X \gamma$
$0.120 \pm 0.021 \pm 0.021$		9	CRAWFORD	92B CLE2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma \gamma$
$0.080 \pm 0.009 \pm 0.007$		10	HEINTZ	92 CSB2 $e^+ e^- \rightarrow \ell^+ \ell^- \gamma \gamma$
<sup>8</sup> LEES 11J reports $[\Gamma(\chi_{b1}(2P) \rightarrow \gamma \mathcal{T}(1S))/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(12.4 \pm 0.3 \pm 0.6) \times 10^{-3}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				
<sup>9</sup> Using $B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = (2.57 \pm 0.07)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \gamma \mathcal{T}(1S)) \times 2 B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = (6.47 \pm 1.12 \pm 0.82) \times 10^{-4}$ and $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = 0.105^{+0.003}_{-0.002} \pm 0.013$ .				
<sup>10</sup> Using $B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = (2.57 \pm 0.07)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (11.5 \pm 0.5 \pm 0.5)\%$ and assuming $e\mu$ universality. Supersedes HEINTZ 91.				

$\Gamma(\pi\pi\chi_{b1}(1P))/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>9.1 \pm 1.3</math> OUR AVERAGE</b>				
$9.2 \pm 1.1 \pm 0.8$	31k	11	LEES	11C BABR $e^+ e^- \rightarrow \pi^+ \pi^- X$
$8.6 \pm 2.3 \pm 2.1$		12	CAWLFIELD	06 CLE3 $\mathcal{T}(3S) \rightarrow 2(\gamma\pi\ell)$
<sup>11</sup> LEES 11C measures $B(\mathcal{T}(3S) \rightarrow \chi_{b1}(2P)X) \times B(\chi_{b1}(2P) \rightarrow \chi_{b1}(1P)\pi^+\pi^-) = (1.16 \pm 0.07 \pm 0.12) \times 10^{-3}$ . We derive the value assuming $B(\mathcal{T}(3S) \rightarrow \chi_{b1}(2P)X) = B(\mathcal{T}(3S) \rightarrow \chi_{b1}(2P)\gamma) = (12.6 \pm 1.2) \times 10^{-2}$ .				
<sup>12</sup> CAWLFIELD 06 quote $\Gamma(\chi_b(2P) \rightarrow \pi\pi\chi_b(1P)) = 0.83 \pm 0.22 \pm 0.08 \pm 0.19$ keV assuming l-spin conservation, no D-wave contribution, $\Gamma(\chi_{b1}(2P)) = 96 \pm 16$ keV, and $\Gamma(\chi_{b2}(2P)) = 138 \pm 19$ keV.				

$\Gamma(D^0 X)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>8.8 \pm 1.5 \pm 0.8</math></b>	2243	13	BRIERE	08 CLEO $\mathcal{T}(3S) \rightarrow \gamma D^0 X$
<sup>13</sup> For $p_{D^0} > 2.5$ GeV/c.				

$\Gamma(\pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.1 \pm 1.0 \pm 0.3</math></b>	30	14	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma \pi^+ \pi^- K^+ K^- \pi^0$
<sup>14</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow \pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(39 \pm 8 \pm 9) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.1 \pm 0.5 \pm 0.1</math></b>	10	15	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ \pi^- K^- K_S^0$
<sup>15</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(14 \pm 5 \pm 3) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}$				$\Gamma_8/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>7.7 \pm 3.1 \pm 0.7</math></b>	15	16	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ \pi^- K^- 2\pi^0$
<sup>16</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(97 \pm 30 \pm 26) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.9 \pm 2.0 \pm 0.5</math></b>	36	17	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- 2\pi^0$
<sup>17</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(74 \pm 16 \pm 19) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.0 \pm 0.4 \pm 0.1</math></b>	12	18	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^-$
<sup>18</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(12 \pm 4 \pm 3) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(2\pi^+ 2\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.5 \pm 1.7 \pm 0.5</math></b>	38	19	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- \pi^0$
<sup>19</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(69 \pm 13 \pm 17) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>9.6 \pm 3.5 \pm 0.9</math></b>	27	20	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- 2\pi^0$
<sup>20</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(121 \pm 29 \pm 33) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(3\pi^+ 2\pi^- K^- K_S^0 \pi^0)/\Gamma_{\text{total}}$				$\Gamma_{13}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.7 \pm 2.5 \pm 0.6</math></b>	17	21	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 3\pi^+ 2\pi^- K^- K_S^0 \pi^0$
<sup>21</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 3\pi^+ 2\pi^- K^- K_S^0 \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(85 \pm 23 \pm 22) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(3\pi^+ 3\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{14}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.2 \pm 0.4 \pm 0.1</math></b>	18	22	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 3\pi^+ 3\pi^-$
<sup>22</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 3\pi^+ 3\pi^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(15 \pm 4 \pm 3) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(3\pi^+ 3\pi^- 2\pi^0)/\Gamma_{\text{total}}$				$\Gamma_{15}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>12 \pm 4 \pm 1</math></b>	44	23	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 3\pi^+ 3\pi^- 2\pi^0$
<sup>23</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 3\pi^+ 3\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(150 \pm 30 \pm 40) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(3\pi^+ 3\pi^- K^+ K^-)/\Gamma_{\text{total}}$				$\Gamma_{16}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.0 \pm 0.7 \pm 0.2</math></b>	16	24	ASNER	08A CLEO $\mathcal{T}(3S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^-$
<sup>24</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 3\pi^+ 3\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P))]$ = $(25 \pm 7 \pm 6) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

# Meson Particle Listings

## $\chi_{b1}(2P)$ , $h_b(2P)$ , $\chi_{b2}(2P)$

$\Gamma(3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{17}/\Gamma</math></span>				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.1\pm2.1\pm0.6</math></b>	25	<sup>25</sup> ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+3\pi^-K^+K^-\pi^0$
<sup>25</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P))]$ = $(77 \pm 17 \pm 21) \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(4\pi^+4\pi^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{18}/\Gamma</math></span>				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.7\pm0.6\pm0.2</math></b>	16	<sup>26</sup> ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 4\pi^+4\pi^-$
<sup>26</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 4\pi^+4\pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P))]$ = $(22 \pm 6 \pm 5) \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

$\Gamma(4\pi^+4\pi^-\pi^0)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{19}/\Gamma</math></span>				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>19\pm7\pm2</math></b>	41	<sup>27</sup> ASNER	08A CLEO	$\Upsilon(3S) \rightarrow \gamma 4\pi^+4\pi^-\pi^0$
<sup>27</sup> ASNER 08A reports $[\Gamma(\chi_{b1}(2P) \rightarrow 4\pi^+4\pi^-\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P))]$ = $(241 \pm 47 \pm 72) \times 10^{-6}$ which we divide by our best value $\text{B}(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P)) = (12.6 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

### $\chi_{b1}(2P)$ Cross-Particle Branching Ratios

$\Gamma(\chi_{b1}(2P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P))/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_3/\Gamma \times \Gamma_{21}^{\Upsilon(3S)}/\Gamma \Upsilon(3S)</math></span>				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>12.4\pm0.3\pm0.6</math></b>	15k	LEES	11J BABR	$\Upsilon(3S) \rightarrow X \gamma$

$\Gamma(\chi_{b1}(2P) \rightarrow \gamma \Upsilon(2S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma\chi_{b1}(2P))/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_2/\Gamma \times \Gamma_{21}^{\Upsilon(3S)}/\Gamma \Upsilon(3S)</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.4\pm0.1\pm0.2</math></b>	4.3k	LEES	11J BABR	$\Upsilon(3S) \rightarrow X \gamma$

$\text{B}(\chi_{b1}(2P) \rightarrow \chi_{b1}(1P)\pi^+\pi^-) \times \text{B}(\Upsilon(3S) \rightarrow \chi_{b1}(2P)X)$				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.16\pm0.07\pm0.12</math></b>	31k	LEES	11c BABR	$e^+e^- \rightarrow \pi^+\pi^-X$

$\text{B}(\chi_{b2}(2P) \rightarrow \rho X + \bar{\rho} X)/\text{B}(\chi_{b1}(2P) \rightarrow \rho X + \bar{\rho} X)$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.109\pm0.007\pm0.040</math></b>	BRIERE	07 CLEO	$\Upsilon(3S) \rightarrow \gamma\chi_{bJ}(2P)$	

$\text{B}(\chi_{b0}(2P) \rightarrow \rho X + \bar{\rho} X)/\text{B}(\chi_{b1}(2P) \rightarrow \rho X + \bar{\rho} X)$				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>1.082\pm0.025\pm0.060</math></b>	BRIERE	07 CLEO	$\Upsilon(3S) \rightarrow \gamma\chi_{bJ}(2P)$	

### $\chi_{b1}(2P)$ REFERENCES

LEES	11C	PR D84 011104	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	11J	PR D84 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ASNER	08A	PR D78 091103	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	08	PR D78 092007	R.A. Briere <i>et al.</i>	(CLEO Collab.)
BRIERE	07	PR D76 012005	R.A. Briere <i>et al.</i>	(CLEO Collab.)
CAWLFIELD	06	PR D73 012003	C. Cawfield <i>et al.</i>	(CLEO Collab.)
ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
CRONIN-HEN...	04	PRL 92 222002	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
CRAWFORD	92B	PL B294 139	G. Crawford, R. Fulton	(CLEO Collab.)
HEINTZ	92	PR D46 1928	U. Heintz <i>et al.</i>	(CUSB II Collab.)
HEINTZ	91	PRL 66 1563	U. Heintz <i>et al.</i>	(CUSB Collab.)
MORRISON	91	PRL 67 1696	R.J. Morrison <i>et al.</i>	(CLEO Collab.)
NARAIN	91	PRL 66 3113	M. Narain <i>et al.</i>	(CUSB Collab.)

## $h_b(2P)$

$$J^G(J^{PC}) = ?^?(1^{+-})$$

OMITTED FROM SUMMARY TABLE

Quantum numbers are quark model predictions.

### $h_b(2P)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>10259.8\pm0.5\pm1.1</math></b>	90k	MIZUK	12 BELL	$e^+e^- \rightarrow \pi^+\pi^-\text{ hadrons}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$10259.8\pm0.6\pm1.4_{-1.0}^{+1.4}$	83.9k	<sup>1</sup> ADACHI	12 BELL	$10.86\,e^+e^- \rightarrow \pi^+\pi^-\text{ MM}$
<sup>1</sup> Superseded by MIZUK 12.				

### $h_b(2P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ hadrons	not seen
$\Gamma_2$ $\eta_b(1S)\gamma$	$(22\pm 5)\%$
$\Gamma_3$ $\eta_b(2S)\gamma$	$(48\pm13)\%$

### $h_b(2P)$ BRANCHING RATIOS

$\Gamma(\text{hadrons})/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_1/\Gamma</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>not seen</b>	83.9k	ADACHI	12 BELL	$10.86\,e^+e^- \rightarrow \pi^+\pi^-\text{ MM}$

$\Gamma(\eta_b(1S)\gamma)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_2/\Gamma</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>22.3\pm3.8\pm3.1_{-3.3}^{+3.1}</math></b>	10k	MIZUK	12 BELL	$e^+e^- \rightarrow (\gamma)\pi^+\pi^-\text{ hadrons}$

$\Gamma(\eta_b(2S)\gamma)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_3/\Gamma</math></span>				
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>47.5\pm10.5\pm6.8_{-7.7}^{+6.8}</math></b>	26k	MIZUK	12 BELL	$e^+e^- \rightarrow (\gamma)\pi^+\pi^-\text{ hadrons}$

### $h_b(2P)$ REFERENCES

ADACHI	12	PRL 108 032001	I. Adachi <i>et al.</i>	(BELLE Collab.)
MIZUK	12	PRL 109 232002	R. Mizuk <i>et al.</i>	(BELLE Collab.)

## $\chi_{b2}(2P)$

$$J^G(J^{PC}) = 0^+(2^{++})$$

$J$  needs confirmation.

Observed in radiative decay of the  $\Upsilon(3S)$ , therefore  $C = +$ . Branching ratio requires E1 transition, M1 is strongly disfavored, therefore  $P = +$ .

### $\chi_{b2}(2P)$ MASS

VALUE (MeV)	DOCUMENT ID
<b><math>10268.65\pm0.22\pm0.50</math> OUR EVALUATION</b>	From $\gamma$ energy below, using $\Upsilon(3S)$ mass = $10355.2 \pm 0.5$ MeV

### $m_{\chi_{b2}(2P)} - m_{\chi_{b1}(2P)}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>13.4\pm0.6</math> OUR AVERAGE</b>			
$12.3\pm2.6\pm0.6$	<sup>1</sup> AAIJ	14B6 LHCb	$p\bar{p} \rightarrow \gamma\mu^+\mu^-X$
$13.5\pm0.4\pm0.5$	<sup>2</sup> HEINTZ	92 CSB2	$e^+e^- \rightarrow \gamma X, \ell^+\ell^- \gamma\gamma$
<sup>1</sup> From the $\chi_{bJ}(2P) \rightarrow \Upsilon(1S)\gamma$ transition.			
<sup>2</sup> From the average photon energy for inclusive and exclusive events. Supersedes NARAIN 91.			

### $\gamma$ ENERGY IN $\Upsilon(3S)$ DECAY

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>86.19\pm0.22</math> OUR EVALUATION</b>				Treating systematic errors as correlated
<b><math>86.40\pm0.18</math> OUR AVERAGE</b>				
$86.04\pm0.06\pm0.27$		ARTUSO	05 CLEO	$\Upsilon(3S) \rightarrow \gamma X$
$86 \pm 1$	101	CRAWFORD	92B CLE2	$e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$
$86.7 \pm 0.4$	10319	<sup>3</sup> HEINTZ	92 CSB2	$e^+e^- \rightarrow \gamma X$
$86.9 \pm 0.4$	157	<sup>4</sup> HEINTZ	92 CSB2	$e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$
$86.4 \pm 0.1 \pm 0.4$	30741	MORRISON	91 CLE2	$e^+e^- \rightarrow \gamma X$
<sup>3</sup> A systematic uncertainty on the energy scale of 0.9% not included. Supersedes NARAIN 91.				
<sup>4</sup> A systematic uncertainty on the energy scale of 0.9% not included. Supersedes HEINTZ 91.				

### $\chi_{b2}(2P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $\omega \Upsilon(1S)$	$(1.10^{+0.34}_{-0.30})\%$	
$\Gamma_2$ $\gamma \Upsilon(2S)$	$(10.6 \pm 2.6)\%$	S=2.0
$\Gamma_3$ $\gamma \Upsilon(1S)$	$(7.0 \pm 0.7)\%$	
$\Gamma_4$ $\pi\pi\chi_{b2}(1P)$	$(5.1 \pm 0.9) \times 10^{-3}$	
$\Gamma_5$ $D^0 X$	$< 2.4\%$	CL=90%
$\Gamma_6$ $\pi^+\pi^-K^+K^-\pi^0$	$< 1.1 \times 10^{-4}$	CL=90%
$\Gamma_7$ $2\pi^+\pi^-K^-K_S^0$	$< 9 \times 10^{-5}$	CL=90%
$\Gamma_8$ $2\pi^+\pi^-K^-K_S^0 2\pi^0$	$< 7 \times 10^{-4}$	CL=90%
$\Gamma_9$ $2\pi^+2\pi^-2\pi^0$	$(3.9 \pm 1.6) \times 10^{-4}$	



See key on page 601

## Meson Particle Listings

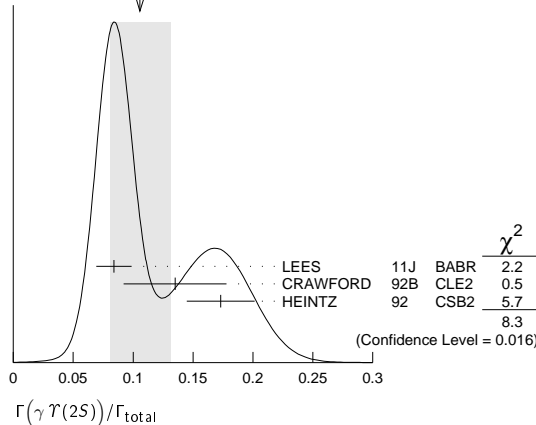
 $\chi_{b2}(2P)$ 

$\Gamma_{10}$	$2\pi^+ 2\pi^- K^+ K^-$	$(9 \pm 4) \times 10^{-5}$	
$\Gamma_{11}$	$2\pi^+ 2\pi^- K^+ K^- \pi^0$	$(2.4 \pm 1.1) \times 10^{-4}$	
$\Gamma_{12}$	$2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	$(4.7 \pm 2.3) \times 10^{-4}$	
$\Gamma_{13}$	$3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	$< 4 \times 10^{-4}$	CL=90%
$\Gamma_{14}$	$3\pi^+ 3\pi^-$	$(9 \pm 4) \times 10^{-5}$	
$\Gamma_{15}$	$3\pi^+ 3\pi^- 2\pi^0$	$(1.2 \pm 0.4) \times 10^{-3}$	
$\Gamma_{16}$	$3\pi^+ 3\pi^- K^+ K^-$	$(1.4 \pm 0.7) \times 10^{-4}$	
$\Gamma_{17}$	$3\pi^+ 3\pi^- K^+ K^- \pi^0$	$(4.2 \pm 1.7) \times 10^{-4}$	
$\Gamma_{18}$	$4\pi^+ 4\pi^-$	$(9 \pm 5) \times 10^{-5}$	
$\Gamma_{19}$	$4\pi^+ 4\pi^- 2\pi^0$	$(1.3 \pm 0.5) \times 10^{-3}$	

 $\chi_{b2}(2P)$  BRANCHING RATIOS

$\Gamma(\omega \mathcal{T}(1S))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE (units $10^{-2}$ )	EVTS
<b><math>1.10^{+0.32+0.11}_{-0.28-0.10}</math> OUR AVERAGE</b>	<b>20.1<math>^{+5.8}_{-5.1}</math></b>
5	CRONIN-HEN..04
TECN	COMMENT
CLE3	$\mathcal{T}(3S) \rightarrow \gamma \omega \mathcal{T}(1S)$
5 Using $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (11.4 \pm 0.8)\%$ and $B(\mathcal{T}(1S) \rightarrow \ell^+ \ell^-) = 2 B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = 2(2.48 \pm 0.06)\%$ .	

$\Gamma(\gamma \mathcal{T}(2S))/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE	EVTS
<b><math>0.106 \pm 0.026</math> OUR AVERAGE</b>	Error includes scale factor of 2.0. See the ideogram below.
0.084 $\pm 0.011 \pm 0.010$	2.5k
0.135 $\pm 0.025 \pm 0.035$	6 LEES
0.173 $\pm 0.021 \pm 0.019$	11J BABR
	92B CLE2
	92 CSB2
	92 HEINTZ
6 LEES 11J reports $[\Gamma(\chi_{b2}(2P) \rightarrow \gamma \mathcal{T}(2S))/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(1.1 \pm 0.1 \pm 0.1) \times 10^{-2}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	
7 Using $B(\mathcal{T}(2S) \rightarrow \mu^+ \mu^-) = (1.37 \pm 0.26)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \gamma \mathcal{T}(2S)) \times 2 B(\mathcal{T}(2S) \rightarrow \mu^+ \mu^-) = (4.98 \pm 0.94 \pm 0.62) \times 10^{-4}$ , and $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = 0.135 \pm 0.003 \pm 0.017$ .	
8 Using $B(\mathcal{T}(2S) \rightarrow \mu^+ \mu^-) = (1.44 \pm 0.10)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (11.1 \pm 0.5 \pm 0.4)\%$ and assuming $e\mu$ universality. Supersedes HEINTZ 91.	

WEIGHTED AVERAGE  
0.106 $\pm 0.026$  (Error scaled by 2.0)

$\Gamma(\gamma \mathcal{T}(1S))/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE	EVTS
<b><math>0.070 \pm 0.007</math> OUR AVERAGE</b>	
0.070 $\pm 0.004 \pm 0.008$	11k
0.072 $\pm 0.014 \pm 0.013$	9 LEES
0.070 $\pm 0.010 \pm 0.006$	10 CRAWFORD
	11 HEINTZ
	92 CSB2
9 LEES 11J reports $[\Gamma(\chi_{b2}(2P) \rightarrow \gamma \mathcal{T}(1S))/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(9.2 \pm 0.3 \pm 0.4) \times 10^{-3}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	
10 Using $B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = (2.57 \pm 0.07)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \gamma \mathcal{T}(2S)) \times 2 B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = (5.03 \pm 0.94 \pm 0.63) \times 10^{-4}$ , and $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = 0.135 \pm 0.003 \pm 0.017$ .	
11 Using $B(\mathcal{T}(1S) \rightarrow \mu^+ \mu^-) = (2.57 \pm 0.07)\%$ , $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (11.1 \pm 0.5 \pm 0.4)\%$ and assuming $e\mu$ universality. Supersedes HEINTZ 91.	

$\Gamma(\pi \pi \chi_{b2}(1P))/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
VALUE (units $10^{-3}$ )	EVTS
<b><math>5.1 \pm 0.9</math> OUR AVERAGE</b>	
4.9 $\pm 0.7 \pm 0.6$	17k
6.0 $\pm 1.6 \pm 1.4$	12 LEES
	13 CRAWFIELD
	11c BABR
	06 CLE3
12 $(0.64 \pm 0.05 \pm 0.08) \times 10^{-3}$ . We derive the value assuming $B(\mathcal{T}(3S) \rightarrow \chi_{b2}(2P) X) = B(\mathcal{T}(3S) \rightarrow \chi_{b2}(2P) \gamma) = (13.1 \pm 1.6) \times 10^{-2}$ .	
13 CRAWFIELD 06 quote $\Gamma(\chi_{b2}(2P) \rightarrow \pi \pi \chi_{b2}(1P)) = 0.83 \pm 0.22 \pm 0.08 \pm 0.19$ keV assuming l-spin conservation, no D-wave contribution, $\Gamma(\chi_{b1}(2P)) = 96 \pm 16$ keV, and $\Gamma(\chi_{b2}(2P)) = 138 \pm 19$ keV.	

$\Gamma(D^0 X)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
VALUE	CL%
<b><math>&lt; 2.4 \times 10^{-2}</math></b>	90
14,15	BRIERE
08	CLEO
$\mathcal{T}(3S) \rightarrow \gamma D^0 X$	

14 For  $p_{D^0} > 2.5$  GeV/c.15 The authors also present their result as  $(0.2 \pm 1.4 \pm 0.1) \times 10^{-2}$ .

$\Gamma(\pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$
VALUE (units $10^{-4}$ )	CL%
<b><math>&lt; 1.1</math></b>	90
16	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma \pi^+ \pi^- K^+ K^- \pi^0$	
16 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow \pi^+ \pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $14 \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = 13.1 \times 10^{-2}$ .	

$\Gamma(2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
VALUE (units $10^{-4}$ )	CL%
<b><math>&lt; 0.9</math></b>	90
17	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ \pi^- K^- K_S^0$	
17 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 2\pi^+ \pi^- K^- K_S^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $12 \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = 13.1 \times 10^{-2}$ .	

$\Gamma(2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
VALUE (units $10^{-4}$ )	CL%
<b><math>&lt; 7</math></b>	90
18	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ \pi^- K^- 2\pi^0$	
18 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 2\pi^+ \pi^- K^- K_S^0 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $87 \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = 13.1 \times 10^{-2}$ .	

$\Gamma(2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>3.9 \pm 1.6 \pm 0.5</math></b>	23
19	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- 2\pi^0$	
19 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 2\pi^+ 2\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(51 \pm 16 \pm 13) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>0.9 \pm 0.4 \pm 0.1</math></b>	11
20	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^-$	
20 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(12 \pm 4 \pm 3) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(2\pi^+ 2\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>2.4 \pm 1.0 \pm 0.3</math></b>	16
21	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- \pi^0$	
21 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(32 \pm 11 \pm 8) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>4.7 \pm 2.2 \pm 0.6</math></b>	14
22	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 2\pi^+ 2\pi^- K^+ K^- 2\pi^0$	
22 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 2\pi^+ 2\pi^- K^+ K^- 2\pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(62 \pm 23 \pm 17) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

$\Gamma(3\pi^+ 2\pi^- K^- K_S^0 \pi^0)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$
VALUE (units $10^{-4}$ )	CL%
<b><math>&lt; 4</math></b>	90
23	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 3\pi^+ 2\pi^- K^- K_S^0 \pi^0$	
23 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 3\pi^+ 2\pi^- K^- K_S^0 \pi^0)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $58 \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = 13.1 \times 10^{-2}$ .	

$\Gamma(3\pi^+ 3\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{14}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS
<b><math>0.9 \pm 0.4 \pm 0.1</math></b>	14
24	ASNER
08A	CLEO
$\mathcal{T}(3S) \rightarrow \gamma 3\pi^+ 3\pi^-$	
24 ASNER 08A reports $[\Gamma(\chi_{b2}(2P) \rightarrow 3\pi^+ 3\pi^-)/\Gamma_{\text{total}}] \times [B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P))]$ = $(12 \pm 4 \pm 3) \times 10^{-6}$ which we divide by our best value $B(\mathcal{T}(3S) \rightarrow \gamma \chi_{b2}(2P)) = (13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.	

# Meson Particle Listings

## $\chi_{b2}(2P)$ , $\Upsilon(3S)$

$\Gamma(3\pi^+3\pi^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{15}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>12\pm 4\pm 1</math></b>	45	<sup>25</sup> ASNER	08A	CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+ 3\pi^- 2\pi^0$

<sup>25</sup> ASNER 08A reports  $[\Gamma(\chi_{b2}(2P) \rightarrow 3\pi^+ 3\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))]$  =  $(159 \pm 33 \pm 43) \times 10^{-6}$  which we divide by our best value  $\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))$  =  $(13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+3\pi^-K^+K^-)/\Gamma_{\text{total}}$					$\Gamma_{16}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.4\pm 0.7\pm 0.2</math></b>	12	<sup>26</sup> ASNER	08A	CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^-$

<sup>26</sup> ASNER 08A reports  $[\Gamma(\chi_{b2}(2P) \rightarrow 3\pi^+ 3\pi^- K^+ K^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))]$  =  $(19 \pm 7 \pm 5) \times 10^{-6}$  which we divide by our best value  $\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))$  =  $(13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(3\pi^+3\pi^-K^+K^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>4.2\pm 1.7\pm 0.5</math></b>	16	<sup>27</sup> ASNER	08A	CLEO	$\Upsilon(3S) \rightarrow \gamma 3\pi^+ 3\pi^- K^+ K^- \pi^0$

<sup>27</sup> ASNER 08A reports  $[\Gamma(\chi_{b2}(2P) \rightarrow 3\pi^+ 3\pi^- K^+ K^- \pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))]$  =  $(55 \pm 16 \pm 15) \times 10^{-6}$  which we divide by our best value  $\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))$  =  $(13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(4\pi^+4\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{18}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.9\pm 0.4\pm 0.1</math></b>	9	<sup>28</sup> ASNER	08A	CLEO	$\Upsilon(3S) \rightarrow \gamma 4\pi^+ 4\pi^-$

<sup>28</sup> ASNER 08A reports  $[\Gamma(\chi_{b2}(2P) \rightarrow 4\pi^+ 4\pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))]$  =  $(12 \pm 5 \pm 3) \times 10^{-6}$  which we divide by our best value  $\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))$  =  $(13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(4\pi^+4\pi^-2\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>13\pm 5\pm 2</math></b>	27	<sup>29</sup> ASNER	08A	CLEO	$\Upsilon(3S) \rightarrow \gamma 4\pi^+ 4\pi^- 2\pi^0$

<sup>29</sup> ASNER 08A reports  $[\Gamma(\chi_{b2}(2P) \rightarrow 4\pi^+ 4\pi^- 2\pi^0)/\Gamma_{\text{total}}] \times [\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))]$  =  $(165 \pm 46 \pm 50) \times 10^{-6}$  which we divide by our best value  $\text{B}(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))$  =  $(13.1 \pm 1.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\chi_{b2}(2P)$ Cross-Particle Branching Ratios

$\Gamma(\chi_{b2}(2P) \rightarrow \gamma \Upsilon(1S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))/\Gamma_{\text{total}}$   
 $\Gamma_3/\Gamma \times \Gamma_{20}^{\Upsilon(3S)}/\Gamma_{\Upsilon(3S)}$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>9.2\pm 0.3\pm 0.4</math></b>	11k	LEES	11J	BABR $\Upsilon(3S) \rightarrow X \gamma$

$\Gamma(\chi_{b2}(2P) \rightarrow \gamma \Upsilon(2S))/\Gamma_{\text{total}} \times \Gamma(\Upsilon(3S) \rightarrow \gamma \chi_{b2}(2P))/\Gamma_{\text{total}}$   
 $\Gamma_2/\Gamma \times \Gamma_{20}^{\Upsilon(3S)}/\Gamma_{\Upsilon(3S)}$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.1\pm 0.1\pm 0.1</math></b>	2.5k	LEES	11J	BABR $\Upsilon(3S) \rightarrow X \gamma$

$\text{B}(\chi_{b2}(2P) \rightarrow \chi_{b2}(1P)\pi^+\pi^-) \times \text{B}(\Upsilon(3S) \rightarrow \chi_{b2}(2P)X)$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.64\pm 0.05\pm 0.08</math></b>	17k	LEES	11c	BABR $e^+e^- \rightarrow \pi^+\pi^-X$

### $\chi_{b2}(2P)$ REFERENCES

AAIJ	14BG	JHEP 1410 088	R. Aaij <i>et al.</i>	(LHCb Collab.)
LEES	11C	PR D84 011104	J.P. Lees <i>et al.</i>	(BABAR Collab.)
LEES	11J	PR D84 072002	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ASNER	08A	PR D78 091103	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BRIERE	08	PR D78 092007	R.A. Briere <i>et al.</i>	(CLEO Collab.)
CRAWFIELD	06	PR D73 012003	C. Crawford <i>et al.</i>	(CLEO Collab.)
ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
CRONIN-HEN.	04	PRL 92 222002	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
CRAWFORD	92B	PL B294 139	G. Crawford, R. Fulton	(CLEO Collab.)
HEINTZ	92	PR D46 1928	U. Heintz <i>et al.</i>	(CUSB II Collab.)
HEINTZ	91	PRL 66 1563	U. Heintz <i>et al.</i>	(CUSB Collab.)
MORRISON	91	PRL 67 1696	R.J. Morrison <i>et al.</i>	(CLEO Collab.)
NARAIN	91	PRL 66 3113	M. Narain <i>et al.</i>	(CUSB Collab.)

## $\Upsilon(3S)$

$I^G(J^{PC}) = 0^-(1^{--})$

### $\Upsilon(3S)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>10355.2\pm 0.5</math></b>	<sup>1</sup> ARTA MONOV 00	MD1	$e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
10355.3 $\pm 0.5$	<sup>2,3</sup> BARU	86B	REDE $e^+e^- \rightarrow \text{hadrons}$

<sup>1</sup> Reanalysis of BARU 86B using new electron mass (COHEN 87).  
<sup>2</sup> Reanalysis of ARTAMONOV 84.  
<sup>3</sup> Superseded by ARTAMONOV 00.

### $m_{\Upsilon(3S)} - m_{\Upsilon(2S)}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>331.50\pm 0.02\pm 0.13</math></b>	LEES	11c	BABR $e^+e^- \rightarrow \pi^+\pi^-X$

### $\Upsilon(3S)$ WIDTH

VALUE (keV)	DOCUMENT ID
<b><math>20.32\pm 1.85</math> OUR EVALUATION</b>	See the Note on "Width Determinations of the $\Upsilon$ States"

### $\Upsilon(3S)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $\Upsilon(2S)$ anything	(10.6 $\pm 0.8$ ) %	
$\Gamma_2$ $\Upsilon(2S)\pi^+\pi^-$	( 2.82 $\pm 0.18$ ) %	S=1.6
$\Gamma_3$ $\Upsilon(2S)\pi^0\pi^0$	( 1.85 $\pm 0.14$ ) %	
$\Gamma_4$ $\Upsilon(2S)\gamma\gamma$	( 5.0 $\pm 0.7$ ) %	
$\Gamma_5$ $\Upsilon(2S)\pi^0$	< 5.1 $\times 10^{-4}$	CL=90%
$\Gamma_6$ $\Upsilon(1S)\pi^+\pi^-$	( 4.37 $\pm 0.08$ ) %	
$\Gamma_7$ $\Upsilon(1S)\pi^0\pi^0$	( 2.20 $\pm 0.13$ ) %	
$\Gamma_8$ $\Upsilon(1S)\eta$	< 1 $\times 10^{-4}$	CL=90%
$\Gamma_9$ $\Upsilon(1S)\pi^0$	< 7 $\times 10^{-5}$	CL=90%
$\Gamma_{10}$ $h_b(1P)\pi^0$	< 1.2 $\times 10^{-3}$	CL=90%
$\Gamma_{11}$ $h_b(1P)\pi^0 \rightarrow \gamma h_b(1S)\pi^0$	( 4.3 $\pm 1.4$ ) $\times 10^{-4}$	
$\Gamma_{12}$ $h_b(1P)\pi^+\pi^-$	< 1.2 $\times 10^{-4}$	CL=90%
$\Gamma_{13}$ $\tau^+\tau^-$	( 2.29 $\pm 0.30$ ) %	
$\Gamma_{14}$ $\mu^+\mu^-$	( 2.18 $\pm 0.21$ ) %	S=2.1
$\Gamma_{15}$ $e^+e^-$	seen	
$\Gamma_{16}$ hadrons		
$\Gamma_{17}$ $ggg$	(35.7 $\pm 2.6$ ) %	
$\Gamma_{18}$ $\gamma\bar{g}g$	( 9.7 $\pm 1.8$ ) $\times 10^{-3}$	
$\Gamma_{19}$ $^2H$ anything	( 2.33 $\pm 0.33$ ) $\times 10^{-5}$	

### Radiative decays

$\Gamma_{20}$ $\gamma\chi_{b2}(2P)$	(13.1 $\pm 1.6$ ) %	S=3.4
$\Gamma_{21}$ $\gamma\chi_{b1}(2P)$	(12.6 $\pm 1.2$ ) %	S=2.4
$\Gamma_{22}$ $\gamma\chi_{b0}(2P)$	( 5.9 $\pm 0.6$ ) %	S=1.4
$\Gamma_{23}$ $\gamma\chi_{b2}(1P)$	( 9.9 $\pm 1.3$ ) $\times 10^{-3}$	S=2.0
$\Gamma_{24}$ $\gamma A^0 \rightarrow \gamma \text{hadrons}$	< 8 $\times 10^{-5}$	CL=90%
$\Gamma_{25}$ $\gamma\chi_{b1}(1P)$	( 9 $\pm 5$ ) $\times 10^{-4}$	S=1.9
$\Gamma_{26}$ $\gamma\chi_{b0}(1P)$	( 2.7 $\pm 0.4$ ) $\times 10^{-3}$	
$\Gamma_{27}$ $\gamma\eta_b(2S)$	< 6.2 $\times 10^{-4}$	CL=90%
$\Gamma_{28}$ $\gamma\eta_b(1S)$	( 5.1 $\pm 0.7$ ) $\times 10^{-4}$	
$\Gamma_{29}$ $\gamma X \rightarrow \gamma + \geq 4$ prongs	[a] < 2.2 $\times 10^{-4}$	CL=95%
$\Gamma_{30}$ $\gamma A^0 \rightarrow \gamma \mu^+\mu^-$	< 5.5 $\times 10^{-6}$	CL=90%
$\Gamma_{31}$ $\gamma A^0 \rightarrow \gamma \tau^+\tau^-$	[b] < 1.6 $\times 10^{-4}$	CL=90%

### Lepton Family number (LF) violating modes

$\Gamma_{32}$ $e^\pm\tau^\mp$	LF	< 4.2 $\times 10^{-6}$	CL=90%
$\Gamma_{33}$ $\mu^\pm\tau^\mp$	LF	< 3.1 $\times 10^{-6}$	CL=90%

[a] 1.5 GeV <  $m_X$  < 5.0 GeV

[b] For  $m_{\tau^+\tau^-}$  in the ranges 4.03–9.52 and 9.61–10.10 GeV.

### $\Upsilon(3S)$ $\Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$

$\Gamma(\text{hadrons}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{16}\Gamma_{15}/\Gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
<b><math>0.414\pm 0.007</math> OUR AVERAGE</b>			
0.413 $\pm 0.004\pm 0.006$	ROSNER	06	CLEO 10.4 $e^+e^- \rightarrow \text{hadrons}$
0.45 $\pm 0.03 \pm 0.03$	<sup>4</sup> GILES	84B	CLEO $e^+e^- \rightarrow \text{hadrons}$

<sup>4</sup> Radiative corrections reevaluated by BUCHMUELLER 88 following KURAEV 85.

# Meson Particle Listings

## $\Upsilon(3S)$

$\Gamma(\Upsilon(1S)\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$					$\Gamma_{615}/\Gamma$
VALUE (eV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>18.46 ± 0.27 ± 0.77</b>	6.4K	<sup>5</sup> AUBERT	08BP BABR	$e^+e^- \rightarrow \gamma\pi^+\pi^-\ell^+\ell^-$	
<sup>5</sup> Using $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$ and $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ .					

$\Upsilon(3S)$ PARTIAL WIDTHS					$\Gamma_{15}$
$\Gamma(e^+e^-)$	VALUE (keV)	DOCUMENT ID			
<b>0.443 ± 0.008 OUR EVALUATION</b>					

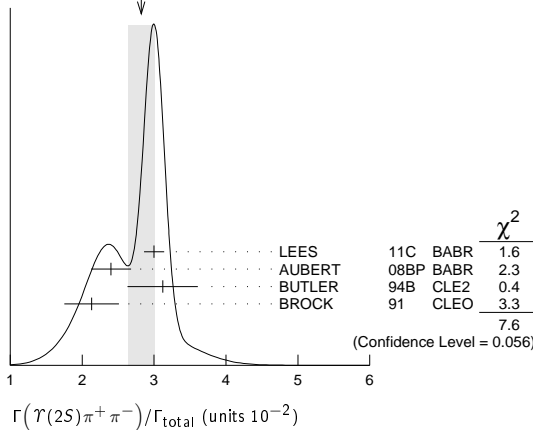
$\Upsilon(3S)$ BRANCHING RATIOS					$\Gamma_1/\Gamma$
$\Gamma(\Upsilon(2S)\text{anything})/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.106 ± 0.008 OUR AVERAGE</b>					
0.1023 ± 0.0105	4625	6,7,8	BUTLER	94B	CLE2 $e^+e^- \rightarrow \ell^+\ell^-X$
0.111 ± 0.012	4891	7,8,9	BROCK	91	CLEO $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-$

- <sup>6</sup> Using  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\gamma\gamma) = (0.038 \pm 0.007)\%$ , and  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^0\pi^0) = (1/2)B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-)$ .
- <sup>7</sup> Using  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.06)\%$ . With the assumption of  $e\mu$  universality.
- <sup>8</sup> Using  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-) = (18.5 \pm 0.8)\%$ .
- <sup>9</sup> Using  $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.31 \pm 0.21)\%$ ,  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\gamma\gamma) \times 2B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (0.188 \pm 0.035)\%$ , and  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^0\pi^0) \times 2B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (0.436 \pm 0.056)\%$ . With the assumption of  $e\mu$  universality.

$\Gamma(\Upsilon(2S)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_2/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.82 ± 0.18 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.	
3.00 ± 0.02 ± 0.14	543k	LEES	11C	BABR $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-$	
2.40 ± 0.10 ± 0.26	800	<sup>10</sup> AUBERT	08BP BABR	$e^+e^- \rightarrow \gamma\pi^+\pi^-\ell^+\ell^-$	
3.12 ± 0.49	980	<sup>11,12</sup> BUTLER	94B	CLE2 $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-$	
2.13 ± 0.38	974	<sup>13</sup> BROCK	91	CLEO $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
4.82 ± 0.65 ± 0.53	138	<sup>13</sup> WU	93	CUSB $\Upsilon(3S) \rightarrow \pi^+\pi^-\ell^+\ell^-$	
3.1 ± 2.0	5	MAGERAS	82	CUSB $\Upsilon(3S) \rightarrow \pi^+\pi^-\ell^+\ell^-$	

- <sup>10</sup> Using  $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$ ,  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ , and  $\Gamma_{ee}(\Upsilon(3S)) = 0.443 \pm 0.008$  keV.
- <sup>11</sup> From the exclusive mode.
- <sup>12</sup> Using  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\gamma\gamma) = (0.038 \pm 0.007)\%$ , and  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^0\pi^0) = (1/2)B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^+\pi^-)$ .
- <sup>13</sup> Using  $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.31 \pm 0.21)\%$ ,  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\gamma\gamma) \times 2B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (0.188 \pm 0.035)\%$ , and  $B(\Upsilon(2S) \rightarrow \Upsilon(1S)\pi^0\pi^0) \times 2B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (0.436 \pm 0.056)\%$ . With the assumption of  $e\mu$  universality.

WEIGHTED AVERAGE  
2.82 ± 0.18 (Error scaled by 1.6)



$\Gamma(\Upsilon(2S)\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_3/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.85 ± 0.14 OUR AVERAGE</b>					
1.82 ± 0.09 ± 0.12	4391	<sup>14</sup> BHARI	09	CLEO $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$	
2.16 ± 0.39		<sup>15,16</sup> BUTLER	94B	CLE2 $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$	
1.7 ± 0.5 ± 0.2	10	<sup>17</sup> HEINTZ	92	CSB2 $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$	

- <sup>14</sup> Authors assume  $B(\Upsilon(1S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.06\%$ .
- <sup>15</sup>  $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.31 \pm 0.21)\%$  and assuming  $e\mu$  universality.
- <sup>16</sup> From the exclusive mode.
- <sup>17</sup>  $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.44 \pm 0.10)\%$  and assuming  $e\mu$  universality. Supersedes HEINTZ 91.

$\Gamma(\Upsilon(2S)\gamma\gamma)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.0502 ± 0.0069</b>	<sup>18</sup> BUTLER	94B	CLE2 $e^+e^- \rightarrow \ell^+\ell^-2\gamma$		
<sup>18</sup> From the exclusive mode.					

$\Gamma(\Upsilon(2S)\pi^0)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.51</b>	90	<sup>19</sup> HE	08A	CLEO $e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$	
<sup>19</sup> Authors assume $B(\Upsilon(2S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.06\%$ .					

$\Gamma(\Upsilon(1S)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$
Abbreviation MM in the COMMENT field below stands for missing mass.					
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.37 ± 0.08 OUR AVERAGE</b>					
4.32 ± 0.07 ± 0.13	90k	<sup>20</sup> LEES	11L	BABR $\Upsilon(3S) \rightarrow \pi^+\pi^-\ell^+\ell^-$	
4.46 ± 0.01 ± 0.13	190k	<sup>21</sup> BHARI	09	CLEO $e^+e^- \rightarrow \pi^+\pi^-$ MM	
4.17 ± 0.06 ± 0.19	6.4K	<sup>22</sup> AUBERT	08BP BABR	$10.58 e^+e^- \rightarrow \gamma\pi^+\pi^-\ell^+\ell^-$	
4.52 ± 0.35	11830	<sup>23</sup> BUTLER	94B	CLE2 $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-$	
4.46 ± 0.34 ± 0.50	451	<sup>23</sup> WU	93	CUSB $\Upsilon(3S) \rightarrow \pi^+\pi^-\ell^+\ell^-$	
4.46 ± 0.30	11221	<sup>23</sup> BROCK	91	CLEO $e^+e^- \rightarrow \pi^+\pi^-\ell^+\ell^-$	

- • • We do not use the following data for averages, fits, limits, etc. • • •
- 4.9 ± 1.0      22      GREEN      82      CLEO       $\Upsilon(3S) \rightarrow \pi^+\pi^-\ell^+\ell^-$
- 3.9 ± 1.3      26      MAGERAS      82      CUSB       $\Upsilon(3S) \rightarrow \pi^+\pi^-\ell^+\ell^-$
- <sup>20</sup> Using  $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$  and  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ .
- <sup>21</sup> A weighted average of the inclusive and exclusive results.
- <sup>22</sup> Using  $B(\Upsilon(2S) \rightarrow e^+e^-) = (1.91 \pm 0.16)\%$ ,  $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.93 \pm 0.17)\%$ , and  $\Gamma_{ee}(\Upsilon(3S)) = 0.443 \pm 0.008$  keV.
- <sup>23</sup> Using  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.06)\%$ . With the assumption of  $e\mu$  universality.

$\Gamma(\Upsilon(2S)\pi^+\pi^-)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$					$\Gamma_2/\Gamma_6$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
0.577 ± 0.026 ± 0.060	800	<sup>24</sup> AUBERT	08BP BABR	$e^+e^- \rightarrow \gamma\pi^+\pi^-\ell^+\ell^-$	
<sup>24</sup> Using $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$ , $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ , $B(\Upsilon(2S) \rightarrow e^+e^-) = (1.91 \pm 0.16)\%$ , and $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.93 \pm 0.17)\%$ . Not independent of other values reported by AUBERT 08BP.					

$\Gamma(\Upsilon(1S)\pi^0\pi^0)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>2.20 ± 0.13 OUR AVERAGE</b>					
2.24 ± 0.09 ± 0.11	6584	<sup>25</sup> BHARI	09	CLEO $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$	
1.99 ± 0.34	56	<sup>26</sup> BUTLER	94B	CLE2 $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$	
2.2 ± 0.4 ± 0.3	33	<sup>27</sup> HEINTZ	92	CSB2 $e^+e^- \rightarrow \pi^0\pi^0\ell^+\ell^-$	
<sup>25</sup> Authors assume $B(\Upsilon(1S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.96\%$ .					
<sup>26</sup> Using $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.06)\%$ and assuming $e\mu$ universality.					
<sup>27</sup> Using $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.57 \pm 0.07)\%$ and assuming $e\mu$ universality. Supersedes HEINTZ 91.					

$\Gamma(\Upsilon(1S)\pi^0\pi^0)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$					$\Gamma_7/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT		
0.501 ± 0.043	<sup>28</sup> BHARI	09	CLEO $e^+e^- \rightarrow \Upsilon(3S)$		
<sup>28</sup> Not independent of other values reported by BHARI 09.					

$\Gamma(\Upsilon(1S)\eta)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.1</b>	90	<sup>29</sup> LEES	11L	BABR $\Upsilon(3S) \rightarrow (\pi^+\pi^-)(\gamma\gamma)\ell^+\ell^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.8	90	<sup>29,30</sup> AUBERT	08BP BABR	$e^+e^- \rightarrow \gamma\pi^+\pi^-\pi^0\ell^+\ell^-$	
<0.18	90	<sup>31</sup> HE	08A	CLEO $e^+e^- \rightarrow \ell^+\ell^-\eta$	
<2.2	90	BROCK	91	CLEO $e^+e^- \rightarrow \ell^+\ell^-\eta$	
<sup>29</sup> Using $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$ , $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ .					
<sup>30</sup> Using $\Gamma_{ee}(\Upsilon(3S)) = 0.443 \pm 0.008$ keV.					
<sup>31</sup> Authors assume $B(\Upsilon(1S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.96\%$ .					

$\Gamma(\Upsilon(1S)\eta)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$					$\Gamma_8/\Gamma_6$
VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.23</b>	90	<sup>32</sup> LEES	11L	BABR $\Upsilon(3S) \rightarrow (\pi^+\pi^-)(\gamma\gamma)\ell^+\ell^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.9	90	<sup>33</sup> AUBERT	08BP BABR	$e^+e^- \rightarrow \gamma\pi^+\pi^-(\pi^0)\ell^+\ell^-$	
<sup>32</sup> Not independent of other values reported by LEES 11L.					
<sup>33</sup> Not independent of other values reported by AUBERT 08BP.					

$\Gamma(\Upsilon(1S)\pi^0)/\Gamma_{\text{total}}$					$\Gamma_9/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.07</b>	90	<sup>34</sup> HE	08A	CLEO $e^+e^- \rightarrow \ell^+\ell^-\gamma\gamma$	
<sup>34</sup> Authors assume $B(\Upsilon(1S) \rightarrow e^+e^-) + B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = 4.96\%$ .					

# Meson Particle Listings

## $\Upsilon(3S)$

### $\Gamma(h_b(1P)\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.2 \times 10^{-3}$	90	35 GE	11 CLEO	$\Upsilon(3S) \rightarrow \pi^0 \text{ anything}$
35 Assuming $M(h_b(1P)) = 9900$ MeV and $\Gamma(h_b(1P)) = 0$ MeV, and allowing $B(h_b(1P) \rightarrow \gamma h_b(1S))$ to vary from 0–100%.				

### $\Gamma(h_b(1P)\pi^0 \rightarrow \gamma \eta_b(1S)\pi^0)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
$4.3 \pm 1.1 \pm 0.9$	LEES	11k BABR	$\Upsilon(3S) \rightarrow \eta_b \gamma \pi^0$

### $\Gamma(h_b(1P)\pi^+\pi^-)/\Gamma_{\text{total}}$ $\Gamma_{12}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
$<1.2$	90	36 LEES	11c BABR	$e^+e^- \rightarrow \pi^+\pi^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<15$		36 BUTLER	94b CLE2	$e^+e^- \rightarrow \pi^+\pi^- X$
$<15$		36 BROCK	91 CLEO	$e^+e^- \rightarrow \pi^+\pi^- X$
36 For $M(h_b(1P)) = 9900$ MeV.				

### $\Gamma(\tau^+\tau^-)/\Gamma_{\text{total}}$ $\Gamma_{13}/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$2.29 \pm 0.21 \pm 0.22$	15k	37 BESSON	07 CLEO	$e^+e^- \rightarrow \Upsilon(3S) \rightarrow \tau^+\tau^-$
37 BESSON 07 reports $[\Gamma(\Upsilon(3S) \rightarrow \tau^+\tau^-)/\Gamma_{\text{total}}] / [B(\Upsilon(3S) \rightarrow \mu^+\mu^-)] = 1.05 \pm 0.08 \pm 0.05$ which we multiply by our best value $B(\Upsilon(3S) \rightarrow \mu^+\mu^-) = (2.18 \pm 0.21) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.				

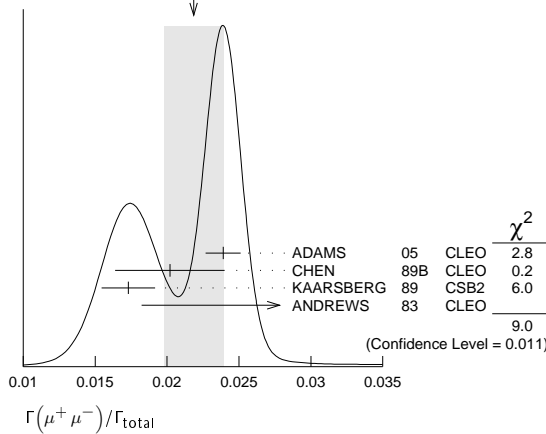
### $\Gamma(\tau^+\tau^-)/\Gamma(\mu^+\mu^-)$ $\Gamma_{13}/\Gamma_{14}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$1.05 \pm 0.08 \pm 0.05$	15k	BESSON	07 CLEO	$e^+e^- \rightarrow \Upsilon(3S)$

### $\Gamma(\mu^+\mu^-)/\Gamma_{\text{total}}$ $\Gamma_{14}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.0218 \pm 0.0021$ OUR AVERAGE				Error includes scale factor of 2.1. See the ideogram below.
$0.0239 \pm 0.0007 \pm 0.0010$	81k	ADAMS	05 CLEO	$e^+e^- \rightarrow \mu^+\mu^-$
$0.0202 \pm 0.0019 \pm 0.0033$		CHEN	89b CLEO	$e^+e^- \rightarrow \mu^+\mu^-$
$0.0173 \pm 0.0015 \pm 0.0011$		KAARSBERG	89 CSB2	$e^+e^- \rightarrow \mu^+\mu^-$
$0.033 \pm 0.013 \pm 0.007$	1096	ANDREWS	83 CLEO	$e^+e^- \rightarrow \mu^+\mu^-$

WEIGHTED AVERAGE  
0.0218±0.0021 (Error scaled by 2.1)



### $\Gamma(ggg)/\Gamma_{\text{total}}$ $\Gamma_{17}/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$35.7 \pm 2.6$	3M	38 BESSON	06a CLEO	$\Upsilon(3S) \rightarrow \text{hadrons}$
38 Calculated using BESSON 06a value of $\Gamma(\gamma g g)/\Gamma(g g g) = (2.72 \pm 0.06 \pm 0.32 \pm 0.37)\%$ and the PDG 08 values of $B(\Upsilon(2S) + \text{anything}) = (10.6 \pm 0.8)\%$ , $B(\pi^+\pi^-\Upsilon(1S)) = (4.40 \pm 0.10)\%$ , $B(\pi^0\pi^0\Upsilon(1S)) = (2.20 \pm 0.13)\%$ , $B(\gamma\chi_{b2}(2P)) = (13.1 \pm 1.6)\%$ , $B(\gamma\chi_{b1}(2P)) = (12.6 \pm 1.2)\%$ , $B(\gamma\chi_{b0}(2P)) = (5.9 \pm 0.6)\%$ , $B(\gamma\chi_{b0}(1P)) = (0.30 \pm 0.11)\%$ , $B(\mu^+\mu^-) = (2.18 \pm 0.21)\%$ , and $R_{\text{hadrons}} = 3.51$ . The statistical error is negligible and the systematic error is partially correlated with $\Gamma(\gamma g g)/\Gamma_{\text{total}}$ BESSON 06a value.				

### $\Gamma(\gamma g g)/\Gamma_{\text{total}}$ $\Gamma_{18}/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$0.97 \pm 0.18$	60k	39 BESSON	06a CLEO	$\Upsilon(3S) \rightarrow \gamma + \text{hadrons}$
39 Calculated using BESSON 06a values of $\Gamma(\gamma g g)/\Gamma(g g g) = (2.72 \pm 0.06 \pm 0.32 \pm 0.37)\%$ and $\Gamma(g g g)/\Gamma_{\text{total}}$ . The statistical error is negligible and the systematic error is partially correlated with $\Gamma(g g g)/\Gamma_{\text{total}}$ BESSON 06a value.				

### $\Gamma(\gamma g g)/\Gamma(g g g)$ $\Gamma_{18}/\Gamma_{17}$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$2.72 \pm 0.06 \pm 0.49$	3M	BESSON	06a CLEO	$\Upsilon(3S) \rightarrow (\gamma +) \text{ hadrons}$

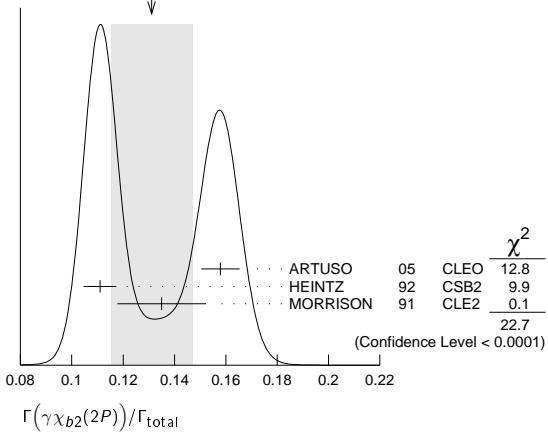
### $\Gamma(2H \text{ anything})/\Gamma_{\text{total}}$ $\Gamma_{19}/\Gamma$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
$2.33 \pm 0.15 \pm 0.31$ $-0.28$	LEES	14g BABR	$e^+e^- \rightarrow 2H X$

### $\Gamma(\gamma\chi_{b2}(2P))/\Gamma_{\text{total}}$ $\Gamma_{20}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.131 \pm 0.016$ OUR AVERAGE				Error includes scale factor of 3.4. See the ideogram below.
$0.1579 \pm 0.0017 \pm 0.0073$	568k	ARTUSO	05 CLEO	$e^+e^- \rightarrow \gamma X$
$0.111 \pm 0.005 \pm 0.004$	10319	40 HEINTZ	92 CSB2	$e^+e^- \rightarrow \gamma X$
$0.135 \pm 0.003 \pm 0.017$	30741	MORRISON	91 CLE2	$e^+e^- \rightarrow \gamma X$
40 Supersedes NARAIN 91.				

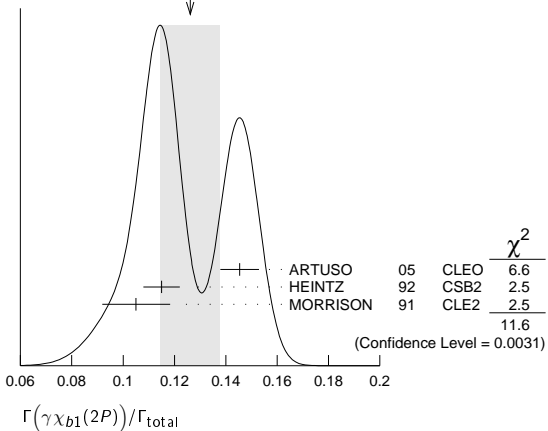
WEIGHTED AVERAGE  
0.131±0.016 (Error scaled by 3.4)



### $\Gamma(\gamma\chi_{b1}(2P))/\Gamma_{\text{total}}$ $\Gamma_{21}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.126 \pm 0.012$ OUR AVERAGE				Error includes scale factor of 2.4. See the ideogram below.
$0.1454 \pm 0.0018 \pm 0.0073$	537k	ARTUSO	05 CLEO	$e^+e^- \rightarrow \gamma X$
$0.115 \pm 0.005 \pm 0.005$	11147	41 HEINTZ	92 CSB2	$e^+e^- \rightarrow \gamma X$
$0.105 \pm 0.003 \pm 0.013$	25759	MORRISON	91 CLE2	$e^+e^- \rightarrow \gamma X$
41 Supersedes NARAIN 91.				

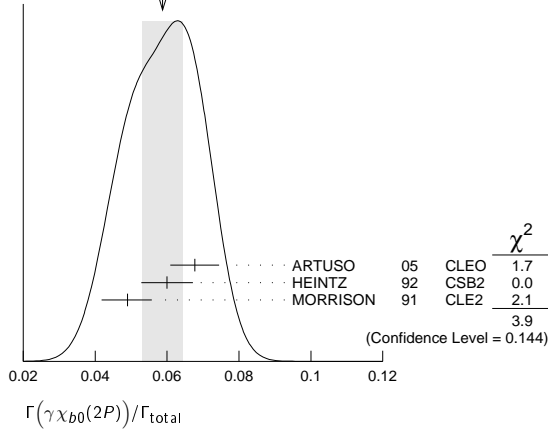
WEIGHTED AVERAGE  
0.126±0.012 (Error scaled by 2.4)



### $\Gamma(\gamma\chi_{b0}(2P))/\Gamma_{\text{total}}$ $\Gamma_{22}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.059 \pm 0.006$ OUR AVERAGE				Error includes scale factor of 1.4. See the ideogram below.
$0.0677 \pm 0.0020 \pm 0.0065$	225k	ARTUSO	05 CLEO	$e^+e^- \rightarrow \gamma X$
$0.060 \pm 0.004 \pm 0.006$	4959	42 HEINTZ	92 CSB2	$e^+e^- \rightarrow \gamma X$
$0.049 \pm 0.003 \pm 0.006$	9903	MORRISON	91 CLE2	$e^+e^- \rightarrow \gamma X$

See key on page 601

Meson Particle Listings  
 $\Upsilon(3S)$ <sup>42</sup> Supersedes NARAIN 91.WEIGHTED AVERAGE  
0.059±0.006 (Error scaled by 1.4) $\Gamma(\Upsilon\chi_{b2}(1P))/\Gamma_{\text{total}}$   $\Gamma_{23}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.9±1.3 OUR AVERAGE</b>			Error includes scale factor of 2.0.		
7.5±1.2±0.5	126	43,44	KORNICER 11	CLEO	$e^+e^- \rightarrow \gamma\Upsilon\ell^+\ell^-$
10.5±0.3±0.7	9.7k		LEES	11J	BABR $\Upsilon(3S) \rightarrow X\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<19	90	45	ASNER	08A	CLEO $\Upsilon(3S) \rightarrow \gamma + \text{hadrons}$
seen		46	HEINTZ	92	CSB2 $e^+e^- \rightarrow \gamma\Upsilon\ell^+\ell^-$

<sup>43</sup> Assuming  $B(\Upsilon(1S) \rightarrow \ell^+\ell^-) = (2.48 \pm 0.05)\%$ .<sup>44</sup> KORNICER 11 reports  $[\Gamma(\Upsilon(3S) \rightarrow \Upsilon\chi_{b2}(1P))/\Gamma_{\text{total}}] \times [B(\chi_{b2}(1P) \rightarrow \gamma\Upsilon(1S))]$  =  $(1.435 \pm 0.162 \pm 0.169) \times 10^{-3}$  which we divide by our best value  $B(\chi_{b2}(1P) \rightarrow \gamma\Upsilon(1S)) = (19.1 \pm 1.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>45</sup> ASNER 08A reports  $[\Gamma(\Upsilon(3S) \rightarrow \Upsilon\chi_{b2}(1P))/\Gamma_{\text{total}}] / [B(\Upsilon(2S) \rightarrow \Upsilon\chi_{b2}(1P))]$  <  $27.1 \times 10^{-2}$  which we multiply by our best value  $B(\Upsilon(2S) \rightarrow \Upsilon\chi_{b2}(1P)) = 7.15 \times 10^{-2}$ .<sup>46</sup> HEINTZ 92, while unable to distinguish between different  $J$  states, measures  $\sum_J B(\Upsilon(3S) \rightarrow \Upsilon\chi_{bJ}) \times B(\chi_{bJ} \rightarrow \gamma\Upsilon(1S)) = (1.7 \pm 0.4 \pm 0.6) \times 10^{-3}$  for  $J = 0, 1, 2$  using inclusive  $\Upsilon(1S)$  decays and  $(1.2^{+0.4}_{-0.3} \pm 0.09) \times 10^{-3}$  for  $J = 1, 2$  using  $\Upsilon(1S) \rightarrow \ell^+\ell^-$ . $\Gamma(\Upsilon\chi_{b1}(1P))/\Gamma_{\text{total}}$   $\Gamma_{25}/\Gamma$ 

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.9±0.5 OUR AVERAGE</b>			Error includes scale factor of 1.9.		
1.6±0.5±0.1	50	47,48	KORNICER 11	CLEO	$e^+e^- \rightarrow \gamma\Upsilon\ell^+\ell^-$
0.5±0.3±0.2			LEES	11J	BABR $\Upsilon(3S) \rightarrow X\gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.7	90	49	ASNER	08A	CLEO $\Upsilon(3S) \rightarrow \gamma + \text{hadrons}$
seen		50	HEINTZ	92	CSB2 $e^+e^- \rightarrow \gamma\Upsilon\ell^+\ell^-$

<sup>47</sup> Assuming  $B(\Upsilon(1S) \rightarrow \ell^+\ell^-) = (2.48 \pm 0.05)\%$ .<sup>48</sup> KORNICER 11 reports  $[\Gamma(\Upsilon(3S) \rightarrow \Upsilon\chi_{b1}(1P))/\Gamma_{\text{total}}] \times [B(\chi_{b1}(1P) \rightarrow \gamma\Upsilon(1S))]$  =  $(5.38 \pm 1.20 \pm 0.95) \times 10^{-4}$  which we divide by our best value  $B(\chi_{b1}(1P) \rightarrow \gamma\Upsilon(1S)) = (33.9 \pm 2.2) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>49</sup> ASNER 08A reports  $[\Gamma(\Upsilon(3S) \rightarrow \Upsilon\chi_{b1}(1P))/\Gamma_{\text{total}}] / [B(\Upsilon(2S) \rightarrow \Upsilon\chi_{b1}(1P))]$  <  $2.5 \times 10^{-2}$  which we multiply by our best value  $B(\Upsilon(2S) \rightarrow \Upsilon\chi_{b1}(1P)) = 6.9 \times 10^{-2}$ .<sup>50</sup> HEINTZ 92, while unable to distinguish between different  $J$  states, measures  $\sum_J B(\Upsilon(3S) \rightarrow \Upsilon\chi_{bJ}) \times B(\chi_{bJ} \rightarrow \gamma\Upsilon(1S)) = (1.7 \pm 0.4 \pm 0.6) \times 10^{-3}$  for  $J = 0, 1, 2$  using inclusive  $\Upsilon(1S)$  decays and  $(1.2^{+0.4}_{-0.3} \pm 0.09) \times 10^{-3}$  for  $J = 1, 2$  using  $\Upsilon(1S) \rightarrow \ell^+\ell^-$ . $\Gamma(\Upsilon\chi_{b0}(1P))/\Gamma_{\text{total}}$   $\Gamma_{26}/\Gamma$ 

VALUE (units $10^{-2}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.27±0.04 OUR AVERAGE</b>					
0.27±0.04±0.02	2.3k		LEES	11J	BABR $\Upsilon(3S) \rightarrow X\gamma$
0.30±0.04±0.10	8.7k		ARTUSO 05	CLEO	$e^+e^- \rightarrow \gamma X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.8	90	51	ASNER	08A	CLEO $\Upsilon(3S) \rightarrow \gamma + \text{hadrons}$
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<sup>51</sup> ASNER 08A reports  $[\Gamma(\Upsilon(3S) \rightarrow \Upsilon\chi_{b0}(1P))/\Gamma_{\text{total}}] / [B(\Upsilon(2S) \rightarrow \Upsilon\chi_{b0}(1P))]$  <  $21.9 \times 10^{-2}$  which we multiply by our best value  $B(\Upsilon(2S) \rightarrow \Upsilon\chi_{b0}(1P)) = 3.8 \times 10^{-2}$ . $\Gamma(\Upsilon\eta_b(2S))/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 6.2</b>	90	ARTUSO 05	CLEO	$e^+e^- \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<19	90	LEES	11J	BABR $\Upsilon(3S) \rightarrow X\gamma$

 $\Gamma(\Upsilon\eta_b(1S))/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.1±0.7 OUR AVERAGE</b>					
7.1±1.8±1.3	2.3±0.5k	52	BONVICINI	10	CLEO $\Upsilon(3S) \rightarrow \gamma X$
4.8±0.5±0.6	19±3k	52	AUBERT	09AQ	BABR $\Upsilon(3S) \rightarrow \gamma X$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<8.5	90		LEES	11J	BABR $\Upsilon(3S) \rightarrow X\gamma$
4.8±0.5±1.2	19±3k	52,53	AUBERT	08v	BABR $\Upsilon(3S) \rightarrow \gamma X$
<4.3	90	54	ARTUSO	05	CLEO $e^+e^- \rightarrow \gamma X$

<sup>52</sup> Assuming  $\Gamma_{\eta_b(1S)} = 10$  MeV.<sup>53</sup> Systematic error re-evaluated by AUBERT 09AQ.<sup>54</sup> Superseded by BONVICINI 10. $\Gamma(\Upsilon X \rightarrow \gamma + \geq 4 \text{ prongs})/\Gamma_{\text{total}}$   
(1.5 GeV <  $m_X$  < 5.0 GeV)

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.2</b>	95	ROSNER	07A	CLEO $e^+e^- \rightarrow \gamma X$

 $\Gamma(\Upsilon a_1^0 \rightarrow \gamma\mu^+\mu^-)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.5</b>	90	55 AUBERT	09z BABR	$e^+e^- \rightarrow \gamma a_1^0 \rightarrow \gamma \mu^+ \mu^-$
<sup>55</sup> For a narrow scalar or pseudoscalar $a_1^0$ with mass in the range 212–9300 MeV, excluding $J/\psi$ and $\psi(2S)$ . Measured 90% CL limits as a function of $m_{a_1^0}$ range from 0.27–5.5 $\times 10^{-6}$ .				

 $\Gamma(\Upsilon a_1^0 \rightarrow \gamma\tau^+\tau^-)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.6 \times 10^{-4}</math></b>	90	<sup>56</sup> AUBERT	09P	BABR $e^+e^- \rightarrow \gamma a_1^0 \rightarrow \gamma \tau^+ \tau^-$
<sup>56</sup> For a narrow scalar or pseudoscalar $a_1^0$ with $M(\tau^+\tau^-)$ in the ranges 4.03–9.52 and 9.61–10.10 GeV. Measured 90% CL limits as a function of $M(\tau^+\tau^-)$ range from $1.5\text{--}16 \times 10^{-5}$ .				

 $\Gamma(\Upsilon A^0 \rightarrow \gamma\text{hadrons})/\Gamma_{\text{total}}$   
(0.3 GeV <  $m_{A^0}$  < 7 GeV)

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;8 \times 10^{-5}</math></b>	90	57 LEES	11H BABR	$\Upsilon(3S) \rightarrow \gamma$ hadrons
57 For a narrow scalar or pseudoscalar $A^0$ , excluding known resonances, with mass in the range 0.3–7 GeV. Measured 90% CL limits as a function of $m_{A^0}$ range from $1 \times 10^{-6}$ to $8 \times 10^{-5}$ .				

## LEPTON FAMILY NUMBER (LF) VIOLATING MODES

 $\Gamma(e^\pm\tau^\mp)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.2</b>	90	LEES	10B	BABR $e^+e^- \rightarrow e^\pm\tau^\mp$

 $\Gamma(\mu^\pm\tau^\mp)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.1</b>	90	LEES	10B	BABR $e^+e^- \rightarrow \mu^\pm\tau^\mp$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<20.3	95	LOVE	08A	CLEO $e^+e^- \rightarrow \mu^\pm\tau^\mp$

 $\Upsilon(3S)$  REFERENCES

LEES	14G	PR D89 111102	J.P. Lees et al.	(BABAR Collab.)
GE	11	PR D84 032008	J.Y. Ge et al.	(CLEO Collab.)
KORNICER	11	PR D83 054003	M. Koricner et al.	(CLEO Collab.)
LEES	11C	PR D84 011104	J.P. Lees et al.	(BABAR Collab.)
LEES	11H	PRL 107 221803	J.P. Lees et al.	(BABAR Collab.)
LEES	11J	PR D84 072002	J.P. Lees et al.	(BABAR Collab.)
LEES	11K	PR D84 091101	J.P. Lees et al.	(BABAR Collab.)
LEES	11L	PR D84 092003	J.P. Lees et al.	(BABAR Collab.)
BONVICINI	10	PR D81 031104	G. Bonvicini et al.	(CLEO Collab.)
LEES	10B	PRL 104 151802	J.P. Lees et al.	(BABAR Collab.)
AUBERT	09AQ	PRL 103 161801	B. Aubert et al.	(BABAR Collab.)
AUBERT	09P	PRL 103 181801	B. Aubert et al.	(BABAR Collab.)
AUBERT	09Z	PRL 103 081803	B. Aubert et al.	(BABAR Collab.)
BHARI	09	PR D79 011103	S.R. Bhari et al.	(CLEO Collab.)
ASNER	08A	PR D78 091103	D.M. Asner et al.	(CLEO Collab.)
AUBERT	08BP	PR D78 112002	B. Aubert et al.	(BABAR Collab.)
AUBERT	08V	PRL 101 071801	B. Aubert et al.	(BABAR Collab.)
HE	08A	PRL 101 192001	Q. He et al.	(CLEO Collab.)
LOVE	08A	PRL 101 201601	W. Love et al.	(CLEO Collab.)
PDG	08	PL B667 1	C. Amsler et al.	(PDG Collab.)
BESSON	07	PRL 98 052002	D. Besson et al.	(CLEO Collab.)
ROSNER	07A	PR D76 117102	J.L. Rosner et al.	(CLEO Collab.)
BESSON	06A	PR D74 012003	D. Besson et al.	(CLEO Collab.)
ROSNER	06	PRL 96 092003	J.L. Rosner et al.	(CLEO Collab.)
ADAMS	05	PRL 94 012001	G.S. Adams et al.	(CLEO Collab.)

# Meson Particle Listings

## $\Upsilon(3S)$ , $\chi_{b1}(3P)$ , $\Upsilon(4S)$

ARTUSO	05	PRL 94 032001	M. Artuso <i>et al.</i>	(CLEO Collab.)
ARTAMONOV	00	PL B474 427	A.S. Artamonov <i>et al.</i>	
BUTLER	94B	PR D49 40	F. Butler <i>et al.</i>	(CLEO Collab.)
WU	93	PL B301 307	Q.W. Wu <i>et al.</i>	(CUSB Collab.)
HEINTZ	92	PR D46 1928	U. Heintz <i>et al.</i>	(CUSB II Collab.)
BROCK	91	PR D43 1448	I.C. Brock <i>et al.</i>	(CLEO Collab.)
HEINTZ	91	PRL 66 1563	U. Heintz <i>et al.</i>	(CUSB Collab.)
MORRISON	91	PRL 67 1696	R.J. Morrison <i>et al.</i>	(CLEO Collab.)
NARAIN	91	PRL 66 3113	M. Narain <i>et al.</i>	(CUSB Collab.)
CHEN	89B	PR D39 3528	W.Y. Chen <i>et al.</i>	(CLEO Collab.)
KAARSBERG	89	PRL 62 2077	T.M. Kaarsberg <i>et al.</i>	(CUSB Collab.)
BUCHMUEL...	88	HE $e^+e^-$ Physics 412	W. Buchmueller, S. Cooper	(HANN, DESY, MIT)
Editors: A. Ali and P. Soeding, World Scientific, Singapore				
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RIS C, NBS)
BARU	86B	ZPHY C32 622 (erratum)	S.E. Baru <i>et al.</i>	(NOVO)
KURAEV	85	SJNP 41 466	E.A. Kurayev, V.S. Fadin	(NOVO)
Translated from YAF 41 733.				
ARTAMONOV	84	PL 137B 272	A.S. Artamonov <i>et al.</i>	(NOVO)
GILES	84B	PR D29 1295	R. Giles <i>et al.</i>	(CLEO Collab.)
ANDREWS	83	PRL 50 807	D.E. Andrews <i>et al.</i>	(CLEO Collab.)
GREEN	82	PRL 49 617	J. Green <i>et al.</i>	(CLEO Collab.)
MAGERAS	82	PL 118B 453	G. Mageras <i>et al.</i>	(COLU, CORN, LSU+)

$\chi_{b1}(3P)$

$$J^G(J^{PC}) = 0^+(1^{++})$$

Observed in the radiative decay to  $\Upsilon(1S, 2S, 3S)$ , therefore  $C = +$ .  
 $J$  needs confirmation.

### $\chi_{b1}(3P)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>10512.1<math>\pm</math> 2.1<math>\pm</math> 0.9</b>	351	<sup>1</sup> AAIJ	14BG LHCb	$pp \rightarrow \gamma \mu^+ \mu^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
10515.7 $^{+2.2+1.5}_{-3.9-2.1}$	169	<sup>2</sup> AAIJ	14BG LHCb	$pp \rightarrow \gamma \mu^+ \mu^- X$
10511.3 $\pm$ 1.7 $\pm$ 2.5	182	<sup>3</sup> AAIJ	14BI LHCb	$pp \rightarrow \gamma \mu^+ \mu^- X$
10530 $\pm$ 5 $\pm$ 9		<sup>4</sup> AAD	12A ATLS	$pp \rightarrow \gamma \mu^+ \mu^- X$
10551 $\pm$ 14 $\pm$ 17		<sup>4</sup> ABAZOV	12Q D0	$p\bar{p} \rightarrow \gamma \mu^+ \mu^- X$

<sup>1</sup> The mass of the  $\chi_{b1}(3P)$  state obtained by combining the results of AAIJ 14BG with that of AAIJ 14BI. The first uncertainty is experimental and the second attributable to the unknown mass splitting, assumed to be  $m_{\chi_{b2}(3P)} - m_{\chi_{b1}(3P)} = 10.5 \pm 1.5$  MeV.

<sup>2</sup> From  $\chi_{b1}(3P) \rightarrow \Upsilon(1S, 2S) \gamma$  transitions assuming  $m_{\chi_{b2}(3P)} - m_{\chi_{b1}(3P)} = 10.5 \pm 1.5$  MeV and allowing for  $\pm 30\%$  variation in the  $\chi_{b2}(3P)$  production rate relative to that of  $\chi_{b1}(3P)$ .

<sup>3</sup> From  $\chi_{b1}(3P) \rightarrow \Upsilon(3S) \gamma$  transition assuming  $m_{\chi_{b2}(3P)} - m_{\chi_{b1}(3P)} = 10.5 \pm 1.5$  MeV.

<sup>4</sup> The mass barycenter of the merged lineshapes from the  $J = 1$  and 2 states.

### $\chi_{b1}(3P)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Upsilon(1S) \gamma$	seen
$\Gamma_2 \quad \Upsilon(2S) \gamma$	seen
$\Gamma_3 \quad \Upsilon(3S) \gamma$	seen

### $\chi_{b1}(3P)$ BRANCHING RATIOS

$\Gamma(\Upsilon(1S)\gamma)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
seen		169	<sup>5</sup> AAIJ	14BG LHCb	$pp \rightarrow \gamma \mu^+ \mu^- X$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
seen			AAD	12A ATLS	$pp \rightarrow \gamma \mu^+ \mu^- X$	
seen			ABAZOV	12Q D0	$p\bar{p} \rightarrow \gamma \mu^+ \mu^- X$	

<sup>5</sup> From  $\chi_{b1}(3P) \rightarrow \Upsilon(1S, 2S) \gamma$  transitions assuming  $m_{\chi_{b2}(3P)} - m_{\chi_{b1}(3P)} = 10.5 \pm 1.5$  MeV and allowing for  $\pm 30\%$  variation in the  $\chi_{b2}(3P)$  production rate relative to that of  $\chi_{b1}(3P)$ .

$\Gamma(\Upsilon(2S)\gamma)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
seen		169	<sup>6</sup> AAIJ	14BG LHCb	$pp \rightarrow \gamma \mu^+ \mu^- X$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
seen			AAD	12A ATLS	$pp \rightarrow \gamma \mu^+ \mu^- X$	

<sup>6</sup> From  $\chi_{b1}(3P) \rightarrow \Upsilon(1S, 2S) \gamma$  transitions assuming  $m_{\chi_{b2}(3P)} - m_{\chi_{b1}(3P)} = 10.5 \pm 1.5$  MeV and allowing for  $\pm 30\%$  variation in the  $\chi_{b2}(3P)$  production rate relative to that of  $\chi_{b1}(3P)$ .

$\Gamma(\Upsilon(3S)\gamma)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
seen		182	AAIJ	14BI LHCb	$pp \rightarrow \gamma \mu^+ \mu^- X$	

### $\chi_{b1}(3P)$ REFERENCES

AAIJ	14BG	JHEP 1410 088	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14BI	EPJ C74 3092	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAD	12A	PRL 108 152001	G. Aad <i>et al.</i>	(ATLAS Collab.)
ABAZOV	12Q	PR D86 031103	V.M. Abazov <i>et al.</i>	(D0 Collab.)

$\Upsilon(4S)$   
or  $\Upsilon(10580)$

$$J^G(J^{PC}) = 0^-(1^{--})$$

### $\Upsilon(4S)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>10579.4<math>\pm</math>1.2 OUR AVERAGE</b>			
10579.3 $\pm$ 0.4 $\pm$ 1.2	AUBERT	05Q BABR	$e^+e^- \rightarrow$ hadrons
10580.0 $\pm$ 3.5	<sup>1</sup> BEBEK	87 CLEO	$e^+e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
10577.4 $\pm$ 1.0	<sup>2</sup> LOVELOCK	85 CUSB	$e^+e^- \rightarrow$ hadrons
<sup>1</sup> Reanalysis of BESSON 85. <sup>2</sup> No systematic error given.			

### $\Upsilon(4S)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>20.5<math>\pm</math>2.5 OUR AVERAGE</b>			
20.7 $\pm$ 1.6 $\pm$ 2.5	AUBERT	05Q BABR	$e^+e^- \rightarrow$ hadrons
20 $\pm$ 2 $\pm$ 4	BESSON	85 CLEO	$e^+e^- \rightarrow$ hadrons
• • • We do not use the following data for averages, fits, limits, etc. • • •			
25 $\pm$ 2.5	LOVELOCK	85 CUSB	$e^+e^- \rightarrow$ hadrons

### $\Upsilon(4S)$ DECAY MODES

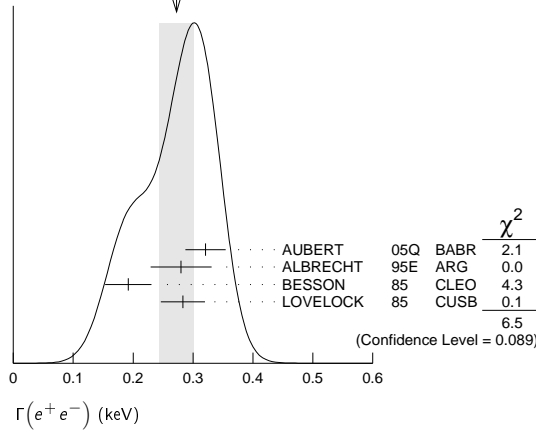
Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \quad B\bar{B}$	$> 96$ %	95%
$\Gamma_2 \quad B^+B^-$	(51.4 $\pm$ 0.6) %	
$\Gamma_3 \quad D^+_s \text{ anything} + \text{ c.c.}$	(17.8 $\pm$ 2.6) %	
$\Gamma_4 \quad B^0\bar{B}^0$	(48.6 $\pm$ 0.6) %	
$\Gamma_5 \quad J/\psi K^0_S + (J/\psi, \eta_c) K^0_S$	$< 4 \times 10^{-7}$	90%
$\Gamma_6 \quad \text{non-} B\bar{B}$	$< 4$ %	95%
$\Gamma_7 \quad e^+e^-$	(1.57 $\pm$ 0.08) $\times 10^{-5}$	
$\Gamma_8 \quad \rho^+\rho^-$	$< 5.7 \times 10^{-6}$	90%
$\Gamma_9 \quad K^*(892)^0\bar{K}^0$	$< 2.0 \times 10^{-6}$	90%
$\Gamma_{10} \quad J/\psi(1S) \text{ anything}$	$< 1.9 \times 10^{-4}$	95%
$\Gamma_{11} \quad D^{*+} \text{ anything} + \text{ c.c.}$	$< 7.4$ %	90%
$\Gamma_{12} \quad \phi \text{ anything}$	(7.1 $\pm$ 0.6) %	
$\Gamma_{13} \quad \phi \eta$	$< 1.8 \times 10^{-6}$	90%
$\Gamma_{14} \quad \phi \eta'$	$< 4.3 \times 10^{-6}$	90%
$\Gamma_{15} \quad \rho \eta$	$< 1.3 \times 10^{-6}$	90%
$\Gamma_{16} \quad \rho \eta'$	$< 2.5 \times 10^{-6}$	90%
$\Gamma_{17} \quad \Upsilon(1S) \text{ anything}$	$< 4 \times 10^{-3}$	
$\Gamma_{18} \quad \Upsilon(1S) \pi^+ \pi^-$	(8.1 $\pm$ 0.6) $\times 10^{-5}$	
$\Gamma_{19} \quad \Upsilon(1S) \eta$	(1.96 $\pm$ 0.28) $\times 10^{-4}$	
$\Gamma_{20} \quad \Upsilon(2S) \pi^+ \pi^-$	(8.6 $\pm$ 1.3) $\times 10^{-5}$	
$\Gamma_{21} \quad h_b(1P) \pi^+ \pi^-$	not seen	
$\Gamma_{22} \quad h_b(1P) \eta$	(2.18 $\pm$ 0.21) $\times 10^{-3}$	
$\Gamma_{23} \quad {}^2H \text{ anything}$	$< 1.3 \times 10^{-5}$	90%

### $\Upsilon(4S)$ PARTIAL WIDTHS

$\Gamma(e^+e^-)$	VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_7$
<b>0.272<math>\pm</math>0.029 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.	
0.321 $\pm$ 0.017 $\pm$ 0.029		AUBERT	05Q BABR	$e^+e^- \rightarrow$ hadrons	
0.28 $\pm$ 0.05 $\pm$ 0.01		<sup>3</sup> ALBRECHT	95E ARG	$e^+e^- \rightarrow$ hadrons	
0.192 $\pm$ 0.007 $\pm$ 0.038		BESSON	85 CLEO	$e^+e^- \rightarrow$ hadrons	
0.283 $\pm$ 0.037		LOVELOCK	85 CUSB	$e^+e^- \rightarrow$ hadrons	

<sup>3</sup> Using LEYAOUANC 77 parametrization of  $\Gamma(s)$ .

See key on page 601

Meson Particle Listings  
 $\Upsilon(4S)$ WEIGHTED AVERAGE  
0.272±0.029 (Error scaled by 1.5) $\Upsilon(4S)$  BRANCHING RATIOS $B\bar{B}$  DECAYS

The ratio of branching fraction to charged and neutral B mesons is often derived assuming isospin invariance in the decays, and relies on the knowledge of the  $B^+/B^0$  lifetime ratio. "OUR EVALUATION" is obtained based on averages of rescaled data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account the common dependence of the measurement on the value of the lifetime ratio.

$\Gamma(B^+B^-)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID
<b>0.514±0.006 OUR EVALUATION</b>	Assuming $B(\Upsilon(4S) \rightarrow B\bar{B}) = 1$

$\Gamma(D_s^+ \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID
<b>0.178±0.021±0.016</b>	4 ARTUSO 05B CLE3 $e^+e^- \rightarrow D_s X$

<sup>4</sup> ARTUSO 05B reports  $[\Gamma(\Upsilon(4S) \rightarrow D_s^+ \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)] = (8.0 \pm 0.2 \pm 0.9) \times 10^{-3}$  which we divide by our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(B^0\bar{B}^0)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID
<b>0.486±0.006 OUR EVALUATION</b>	Assuming $B(\Upsilon(4S) \rightarrow B\bar{B}) = 1$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.487±0.010±0.008	<sup>5</sup> AUBERT,B 05H BABR $\Upsilon(4S) \rightarrow D^*\ell\nu_\ell$
<sup>5</sup> Direct measurement. This value is averaged with the value extracted from the $\Gamma(B^+B^-)/\Gamma(B^0\bar{B}^0)$ measurements.	

$\Gamma(B^+B^-)/\Gamma(B^0\bar{B}^0)$	$\Gamma_2/\Gamma_4$
VALUE	DOCUMENT ID
<b>1.058±0.024 OUR EVALUATION</b>	

1.006±0.036±0.031	<sup>6</sup> AUBERT 04F BABR $\Upsilon(4S) \rightarrow B\bar{B} \rightarrow J/\psi K$
1.01 ± 0.03 ± 0.09	<sup>6</sup> HASTINGS 03 BELL $\Upsilon(4S) \rightarrow B\bar{B} \rightarrow \text{dileptons}$
1.058±0.084±0.136	<sup>7</sup> ATHAR 02 CLEO $\Upsilon(4S) \rightarrow B\bar{B} \rightarrow D^*\ell\nu$
1.10 ± 0.06 ± 0.05	<sup>8</sup> AUBERT 02 BABR $\Upsilon(4S) \rightarrow B\bar{B} \rightarrow (c\bar{c})K^*$
1.04 ± 0.07 ± 0.04	<sup>9</sup> ALEXANDER 01 CLEO $\Upsilon(4S) \rightarrow B\bar{B} \rightarrow J/\psi K^*$

<sup>6</sup> HASTINGS 03 and AUBERT 04F assume  $\tau(B^+)/\tau(B^0) = 1.083 \pm 0.017$ .

<sup>7</sup> ATHAR 02 assumes  $\tau(B^+)/\tau(B^0) = 1.074 \pm 0.028$ . Supersedes BARISH 95.

<sup>8</sup> AUBERT 02 assumes  $\tau(B^+)/\tau(B^0) = 1.062 \pm 0.029$ .

<sup>9</sup> ALEXANDER 01 assumes  $\tau(B^+)/\tau(B^0) = 1.066 \pm 0.024$ .

$[\Gamma(J/\psi K_S^0) + \Gamma(J/\psi, \eta_c K_S^0)]/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b>&lt;4</b>	<sup>10</sup> TAJIMA 07A BELL $\Upsilon(4S) \rightarrow B^0\bar{B}^0$

<sup>10</sup>  $\Upsilon(4S)$  with  $CP = +1$  decays to the final state with  $CP = -1$ .

non- $B\bar{B}$  DECAYS

$\Gamma(\text{non-}B\bar{B})/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID
<b>&lt;0.04</b>	95 BARISH 96B CLEO $e^+e^-$

$\Gamma(e^+e^-)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID
<b>1.57±0.08 OUR AVERAGE</b>	

1.55±0.04±0.07	AUBERT 05Q BABR $e^+e^- \rightarrow \text{hadrons}$
2.77±0.50±0.49	<sup>11</sup> ALBRECHT 95E ARG $e^+e^- \rightarrow \text{hadrons}$

<sup>11</sup> Using LEYAOUANC 77 parametrization of  $\Gamma(s)$ .

$\Gamma(\rho^+\rho^-)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$
VALUE	DOCUMENT ID
<b>&lt;5.7 × 10<sup>-6</sup></b>	90 AUBERT 08B BABR $e^+e^- \rightarrow \pi^+\pi^-2\pi^0$

$\Gamma(K^*(892)^0\bar{K}^0)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$
VALUE	DOCUMENT ID
<b>&lt;2.0 × 10<sup>-6</sup></b>	90 SHEN 13A BELL $e^+e^- \rightarrow K^*(892)^0\bar{K}^0$

$\Gamma(J/\psi(1S) \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>&lt;1.9</b>	95 <sup>12</sup> ABE 02D BELL $e^+e^- \rightarrow J/\psi X \rightarrow \ell^+\ell^-X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<4.7 90 <sup>12</sup> AUBERT 01C BABR  $e^+e^- \rightarrow J/\psi X \rightarrow \ell^+\ell^-X$

<sup>12</sup> Uses  $B(J/\psi \rightarrow e^+e^-) = 0.0593 \pm 0.0010$  and  $B(J/\psi \rightarrow \mu^+\mu^-) = 0.0588 \pm 0.0010$ .

$\Gamma(D^{*+} \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$
VALUE	DOCUMENT ID
<b>&lt;0.074</b>	90 <sup>13</sup> ALEXANDER 90C CLEO $e^+e^-$

<sup>13</sup> For  $x > 0.473$ .

$\Gamma(\phi \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$
VALUE (units $10^{-2}$ )	DOCUMENT ID
<b>7.1 ± 0.1 ± 0.6</b>	HUANG 07 CLEO $\Upsilon(4S) \rightarrow \phi X$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.23 90 <sup>14</sup> ALEXANDER 90C CLEO  $e^+e^-$

<sup>14</sup> For  $x > 0.52$ .

$\Gamma(\phi\eta)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID
<b>&lt;1.8</b>	90 <sup>15</sup> BELOUS 09 BELL $e^+e^- \rightarrow \phi\eta$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<2.5 90 AUBERT,BE 06F BABR  $e^+e^- \rightarrow \phi\eta$

<sup>15</sup> Using all intermediate branching fraction values from PDG 08.

$\Gamma(\phi\eta')/\Gamma_{\text{total}}$	$\Gamma_{14}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID
<b>&lt;4.3</b>	90 <sup>16</sup> BELOUS 09 BELL $e^+e^- \rightarrow \phi\eta'$

<sup>16</sup> Using all intermediate branching fraction values from PDG 08.

$\Gamma(\rho\eta)/\Gamma_{\text{total}}$	$\Gamma_{15}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID
<b>&lt;1.3</b>	90 <sup>17</sup> BELOUS 09 BELL $e^+e^- \rightarrow \rho\eta$

<sup>17</sup> Using all intermediate branching fraction values from PDG 08.

$\Gamma(\rho\eta')/\Gamma_{\text{total}}$	$\Gamma_{16}/\Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID
<b>&lt;2.5</b>	90 <sup>18</sup> BELOUS 09 BELL $e^+e^- \rightarrow \rho\eta'$

<sup>18</sup> Using all intermediate branching fraction values from PDG 08.

$\Gamma(\Upsilon(1S) \text{ anything})/\Gamma_{\text{total}}$	$\Gamma_{17}/\Gamma$
VALUE	DOCUMENT ID
<b>&lt;0.004</b>	90 ALEXANDER 90C CLEO $e^+e^-$

$\Gamma(\Upsilon(1S)\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_{18}/\Gamma$
VALUE (units $10^{-5}$ )	DOCUMENT ID
<b>8.1 ± 0.6 OUR AVERAGE</b>	

8.5 ± 1.3 ± 0.2	<sup>19</sup> SOKOLOV 09 BELL $e^+e^- \rightarrow \pi^+\pi^-\mu^+\mu^-$
8.00±0.64±0.27	<sup>20</sup> AUBERT 08BP BABR $\Upsilon(4S) \rightarrow \pi^+\pi^-\ell^+\ell^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

17.8 ± 4.0 ± 0.3	<sup>21,22</sup> SOKOLOV 07 BELL $e^+e^- \rightarrow \pi^+\pi^-\mu^+\mu^-$
9.0 ± 1.5 ± 0.2	<sup>23</sup> AUBERT 06R BABR $e^+e^- \rightarrow \pi^+\pi^-\mu^+\mu^-$

<12 90 GLENN 99 CLE2  $e^+e^-$

<sup>19</sup> SOKOLOV 09 reports  $[\Gamma(\Upsilon(4S) \rightarrow \Upsilon(1S)\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(1S) \rightarrow \mu^+\mu^-)] = (0.211 \pm 0.030 \pm 0.014) \times 10^{-5}$  which we divide by our best value  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>20</sup> Using  $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$  and  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ .

<sup>21</sup> SOKOLOV 07 reports  $[\Gamma(\Upsilon(4S) \rightarrow \Upsilon(1S)\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(1S) \rightarrow \mu^+\mu^-)] = (4.42 \pm 0.81 \pm 0.56) \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>22</sup> According to the authors, systematic errors were underestimated.

<sup>23</sup> Superseded by AUBERT 08BP. AUBERT 06R reports  $[\Gamma(\Upsilon(4S) \rightarrow \Upsilon(1S)\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(1S) \rightarrow \mu^+\mu^-)] = (2.23 \pm 0.25 \pm 0.27) \times 10^{-6}$  which we divide by our best value  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

# Meson Particle Listings

## $\Upsilon(4S), X(10610)^{\pm}$

$\Gamma(\Upsilon(1S)\eta)/\Gamma_{\text{total}}$					$\Gamma_{19}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.96\pm 0.26\pm 0.09</math></b>		56	<sup>24</sup> AUBERT	08BP BABR	$\Upsilon(4S) \rightarrow \pi^+\pi^-\pi^0\ell^+\ell^-$	

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<2.7 90 <sup>25</sup>TAMPONI 15 BELL  $e^+e^- \rightarrow \gamma\eta + \text{hadrons}$   
<sup>24</sup> Using  $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$  and  $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ .  
<sup>25</sup> Using  $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20)\%$ .

$\Gamma(\Upsilon(1S)\eta)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$					$\Gamma_{19}/\Gamma_{18}$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$2.41\pm 0.40\pm 0.12$	56	<sup>26</sup> AUBERT	08BP BABR	$\Upsilon(4S) \rightarrow \pi^+\pi^-(\pi^0)\ell^+\ell^-$		
<sup>26</sup> Not independent of other values reported by AUBERT 08BP.						

$\Gamma(\Upsilon(2S)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{20}/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>0.86\pm 0.11\pm 0.07</math></b>		220	<sup>27</sup> AUBERT	08BP BABR	$\Upsilon(4S) \rightarrow \pi^+\pi^-\ell^+\ell^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.88\pm 0.17\pm 0.08$	97 $\pm$ 15	<sup>28</sup> AUBERT	06R BABR	$e^+e^- \rightarrow \pi^+\pi^-\mu^+\mu^-$		
<3.9	90	GLENN	99 CLE2	$e^+e^-$		
<sup>27</sup> Using $B(\Upsilon(2S) \rightarrow e^+e^-) = (1.91 \pm 0.16)\%$ and $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.93 \pm 0.17)\%$ . <sup>28</sup> Superseded by AUBERT 08BP. AUBERT 06R reports $[\Gamma(\Upsilon(4S) \rightarrow \Upsilon(2S)\pi^+\pi^-)/\Gamma_{\text{total}}] \times [B(\Upsilon(2S) \rightarrow \mu^+\mu^-)] = (1.69 \pm 0.26 \pm 0.20) \times 10^{-6}$ which we divide by our best value $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.93 \pm 0.17) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.						

$\Gamma(\Upsilon(2S)\pi^+\pi^-)/\Gamma(\Upsilon(1S)\pi^+\pi^-)$					$\Gamma_{20}/\Gamma_{18}$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$1.16\pm 0.16\pm 0.14$	220	<sup>29</sup> AUBERT	08BP BABR	$\Upsilon(4S) \rightarrow \pi^+\pi^-\ell^+\ell^-$		
<sup>29</sup> Using $B(\Upsilon(1S) \rightarrow e^+e^-) = (2.38 \pm 0.11)\%$ , $B(\Upsilon(1S) \rightarrow \mu^+\mu^-) = (2.48 \pm 0.05)\%$ , $B(\Upsilon(2S) \rightarrow e^+e^-) = (1.91 \pm 0.16)\%$ , and $B(\Upsilon(2S) \rightarrow \mu^+\mu^-) = (1.93 \pm 0.17)\%$ . Not independent of other values reported by AUBERT 08BP.						

$\Gamma(h_b(1P)\pi^+\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{21}/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
not seen	$(35^{+32}_{-26})\text{k}$	<sup>30</sup> ADACHI	12 BELL	$10.58 e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$		
<sup>30</sup> From the upper limit on the ratio of $\sigma(e^+e^- \rightarrow h_b(1P)\pi^+\pi^-)$ at the $\Upsilon(4S)$ to that at the $\Upsilon(5S)$ of 0.27.						

$\Gamma(h_b(1P)\eta)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b><math>2.18\pm 0.11\pm 0.18</math></b>	112k	<sup>31</sup> TAMPONI	15 BELL	$e^+e^- \rightarrow h_b(1P)\eta$		
<sup>31</sup> Using $B(\eta \rightarrow 2\gamma) = (39.41 \pm 0.20)\%$ .						

$\Gamma(\overline{2H} \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$	
VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;1.3</b>	90	ASNER	07 CLEO	$e^+e^- \rightarrow \overline{d}X$		

### $\Upsilon(4S)$ REFERENCES

TAMPONI	15	PRL 115 142001	U. Tamponi <i>et al.</i>	(BELLE Collab.)
SHEN	13A	PR D88 052019	C.P. Shen <i>et al.</i>	(BELLE Collab.)
ADACHI	12	PRL 108 032001	I. Adachi <i>et al.</i>	(BELLE Collab.)
BELOUS	09	PL B681 400	K. Belous <i>et al.</i>	(BELLE Collab.)
SOKOLOV	09	PR D79 051103	A. Sokolov <i>et al.</i>	(BELLE Collab.)
AUBERT	08BO	PR D78 071103	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	08BP	PR D78 112002	B. Aubert <i>et al.</i>	(BABAR Collab.)
PDG	08	PL B667 1	C. Amsler <i>et al.</i>	(PDG Collab.)
ASNER	07	PR D75 012009	D.M. Asner <i>et al.</i>	(CLEO Collab.)
HUANG	07	PR D75 012002	G.S. Huang <i>et al.</i>	(CLEO Collab.)
SOKOLOV	07	PR D75 071103	A. Sokolov <i>et al.</i>	(BELLE Collab.)
TAJIMA	07A	PRL 99 211601	O. Tajima <i>et al.</i>	(BELLE Collab.)
AUBERT	06R	PRL 96 232001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,BE	06F	PR D74 111103	B. Aubert <i>et al.</i>	(BABAR Collab.)
ARTUSO	05B	PRL 95 261801	M. Artuso <i>et al.</i>	(CLEO Collab.)
AUBERT	05Q	PR D72 032005	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT,B	05H	PRL 95 042001	B. Aubert <i>et al.</i>	(BABAR Collab.)
AUBERT	04F	PR D69 071101	B.Aubert <i>et al.</i>	(BABAR Collab.)
HASTINGS	03	PR D67 052004	N.C. Hastings <i>et al.</i>	(BELLE Collab.)
ABE	02D	PRL 88 052001	K. Abe <i>et al.</i>	(BELLE Collab.)
ATHAR	02	PR D66 052003	S.B. Athar <i>et al.</i>	(CLEO Collab.)
AUBERT	02	PR D65 032001	B. Aubert <i>et al.</i>	(BABAR Collab.)
ALEXANDER	01	PRL 86 2737	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
AUBERT	01C	PRL 87 162002	B. Aubert <i>et al.</i>	(BABAR Collab.)
GLENN	99	PR D59 052003	S. Glenn <i>et al.</i>	(CLEO Collab.)
BARISH	96B	PRL 76 1570	B.C. Barish <i>et al.</i>	(CLEO Collab.)
ALBRECHT	95E	ZPHY C65 619	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARISH	95	PR D51 1014	B.C. Barish <i>et al.</i>	(CLEO Collab.)
ALEXANDER	90C	PRL 64 2236	J. Alexander <i>et al.</i>	(CLEO Collab.)
BEBEK	87	PR D36 1289	C. Bebek <i>et al.</i>	(CLEO Collab.)
BESSON	85	PRL 54 381	D. Besson <i>et al.</i>	(CLEO Collab.)
LOVELOCK	85	PRL 54 377	D.M.J. Lovelock <i>et al.</i>	(CUSB Collab.)
LEYAOUANC	77	PL B71 397	A. Le Yaouanc <i>et al.</i>	(ORSAY)

$X(10610)^{\pm}$

$I^G(J^P) = 1^+(1^+)$

Observed by BONDAR 12 in  $\Upsilon(5S)$  decays to  $\Upsilon(nS)\pi^+\pi^-$  (n = 1, 2, 3) and  $h_b(mP)\pi^+\pi^-$  (m = 1, 2).  $J^P = 1^+$  is favored from angular analyses. Isospin = 1 is favored due to observation by KROKOVNY 13 of a corresponding neutral state produced in  $\Upsilon(10860) \rightarrow \Upsilon(2S)/\Upsilon(3S)\pi^0\pi^0$  decays at a consistent mass.

$X(10610)^{\pm}$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b><math>10607.2\pm 2.0</math></b>	<sup>1</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \text{hadrons}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$10608.5\pm 3.4^{+3.7}_{-1.4}$	<sup>2</sup> GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
$10608.1\pm 1.2^{+1.5}_{-0.2}$	<sup>2</sup> GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
$10607.4\pm 1.5^{+0.8}_{-0.2}$	<sup>2</sup> GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
$10611 \pm 4 \pm 3$	<sup>3</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
$10609 \pm 2 \pm 3$	<sup>3</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
$10608 \pm 2 \pm 3$	<sup>3</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
$10605 \pm 2 \pm 3$	<sup>3</sup> BONDAR	12 BELL	$e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$	
$10599 \pm 6 \pm 5$	<sup>3</sup> BONDAR	12 BELL	$e^+e^- \rightarrow h_b(2P)\pi^+\pi^-$	
<sup>1</sup> Average of the BONDAR 12 measurements in separate channels. <sup>2</sup> Correlated with the corresponding result from BONDAR 12. <sup>3</sup> Superseded by the average measurement of BONDAR 12.				

$X(10610)^{\pm}$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b><math>18.4\pm 2.4</math></b>	<sup>4</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \text{hadrons}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$18.5\pm 5.3^{+6.1}_{-2.3}$	<sup>5</sup> GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
$20.8\pm 2.5^{+0.3}_{-2.1}$	<sup>5</sup> GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
$18.7\pm 3.4^{+2.5}_{-1.3}$	<sup>5</sup> GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
$22.3\pm 7.7^{+3.0}_{-4.0}$	<sup>6</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
$24.2\pm 3.1^{+2.0}_{-3.0}$	<sup>6</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
$17.6\pm 3.0\pm 3.0$	<sup>6</sup> BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
$11.4\pm 4.5^{+2.1}_{-3.9}\pm 1.2$	<sup>6</sup> BONDAR	12 BELL	$e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$	
$13 \pm 10 \pm 9$ $-8 -7$	<sup>6</sup> BONDAR	12 BELL	$e^+e^- \rightarrow h_b(2P)\pi^+\pi^-$	
<sup>4</sup> Average of the BONDAR 12 measurements in separate channels. <sup>5</sup> Correlated with the corresponding result from BONDAR 12. <sup>6</sup> Superseded by the average measurement of BONDAR 12.				

$X(10610)^+$ DECAY MODES				
$X(10610)^-$ decay modes are charge conjugates of the modes below.				
Mode		Fraction ( $\Gamma_i/\Gamma$ )		
$\Gamma_1$	$\Upsilon(1S)\pi^+$	seen		
$\Gamma_2$	$\Upsilon(2S)\pi^+$	seen		
$\Gamma_3$	$\Upsilon(3S)\pi^+$	seen		
$\Gamma_4$	$h_b(1P)\pi^+$	seen		
$\Gamma_5$	$h_b(2P)\pi^+$	seen		

$X(10610)^{\pm}$ BRANCHING RATIOS				
$\Gamma(\Upsilon(1S)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
seen	BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
$\Gamma(\Upsilon(2S)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
seen	BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
$\Gamma(\Upsilon(3S)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	GARMASH	15 BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
seen	BONDAR	12 BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	



See key on page 601

Meson Particle Listings

$X(10610)^{\pm}, X(10610)^0, X(10650)^{\pm}$

$\Gamma(h_b(1P)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$	
$\Gamma(h_b(2P)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow h_b(2P)\pi^+\pi^-$	

$X(10610)^{\pm}$  REFERENCES

GARMASH 15	PR D91 072003	A. Garmash <i>et al.</i>	(BELLE Collab.)
KROKOVNY 13	PR D88 052016	P. Krokovny <i>et al.</i>	(BELLE Collab.)
BONDAR 12	PRL 108 122001	A. Bondar <i>et al.</i>	(BELLE Collab.)

$X(10610)^0$

$J^G(J^P) = 1^+(1^+)$

Observed by KROKOVNY 13 in  $\Upsilon(10860) \rightarrow \Upsilon(nS)\pi^0\pi^0$  ( $n=2,3$ ). Isospin 1 is favored from the proximity in mass to  $X(10610)^{\pm}$  and their similarity of observed decay modes and cross sections.  $J^P = 1^+$  is favored from angular analysis of  $X(10610)^{\pm}$  decays by BONDAR 12.

$X(10610)^0$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$10609 \pm 4 \pm 4$	<sup>1</sup> KROKOVNY 13	BELL	$e^+e^- \rightarrow \Upsilon(2S)/\Upsilon(3S)\pi^0\pi^0$
<sup>1</sup> From a simultaneous fit to the KROKOVNY 13 Dalitz analysis of $e^+e^- \rightarrow \Upsilon(2S)/\Upsilon(3S)\pi^0\pi^0$ decays with fixed width $\Gamma(X(10610)^0) = 18.4$ MeV.			

$X(10610)^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Upsilon(1S)\pi^0$	not seen
$\Gamma_2 \quad \Upsilon(2S)\pi^0$	seen
$\Gamma_3 \quad \Upsilon(3S)\pi^0$	seen

$X(10610)^0$  BRANCHING RATIOS

$\Gamma(\Upsilon(1S)\pi^0)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
not seen	KROKOVNY 13	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^0\pi^0$	
$\Gamma(\Upsilon(2S)\pi^0)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	<sup>2</sup> KROKOVNY 13	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^0\pi^0$	
<sup>2</sup> Combined significance in $e^+e^- \rightarrow \Upsilon(2S)/\Upsilon(3S)\pi^0\pi^0$ , including systematics, of 6.5 $\sigma$ .				
$\Gamma(\Upsilon(3S)\pi^0)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	<sup>3</sup> KROKOVNY 13	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^0\pi^0$	
<sup>3</sup> Combined significance in $e^+e^- \rightarrow \Upsilon(2S)/\Upsilon(3S)\pi^0\pi^0$ , including systematics, of 6.5 $\sigma$ .				

$X(10610)^0$  REFERENCES

KROKOVNY 13	PR D88 052016	P. Krokovny <i>et al.</i>	(BELLE Collab.)
BONDAR 12	PRL 108 122001	A. Bondar <i>et al.</i>	(BELLE Collab.)

$X(10650)^{\pm}$

$J^G(J^P) = ?^+(1^+)$

OMITTED FROM SUMMARY TABLE

Observed by BONDAR 12 in  $\Upsilon(5S)$  decays to  $\Upsilon(nS)\pi^+\pi^-$  ( $n = 1, 2, 3$ ) and  $h_b(mP)\pi^+\pi^-$  ( $m = 1, 2$ ).  $J^P = 1^+$  is favored from angular analyses.

$X(10650)^{\pm}$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$10652.2 \pm 1.5$	<sup>1</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$10656.7 \pm 5.0^{+1.1}_{-3.1}$	<sup>2</sup> GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$
$10650.7 \pm 1.5^{+0.5}_{-0.2}$	<sup>2</sup> GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$

$10651.2 \pm 1.0^{+0.4}_{-0.3}$	<sup>2</sup> GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$
$10657 \pm 6 \pm 3$	<sup>3</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$
$10651 \pm 2 \pm 3$	<sup>3</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$
$10652 \pm 1 \pm 2$	<sup>3</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$
$10654 \pm 3 \pm 1_{-2}$	<sup>3</sup> BONDAR 12	BELL	$e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$
$10651 \pm 2_{-3}^{+3}_{-2}$	<sup>3</sup> BONDAR 12	BELL	$e^+e^- \rightarrow h_b(2P)\pi^+\pi^-$

- <sup>1</sup> Average of the BONDAR 12 measurements in separate channels.
- <sup>2</sup> Correlated with the corresponding result from BONDAR 12.
- <sup>3</sup> Superseded by the average measurement of BONDAR 12.

$X(10650)^{\pm}$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$11.5 \pm 2.2$	<sup>4</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$12.1^{+11.3}_{-4.8} \pm 2.7_{-0.6}$	<sup>5</sup> GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$
$14.2 \pm 3.7^{+0.9}_{-0.4}$	<sup>5</sup> GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$
$9.3 \pm 2.2^{+0.3}_{-0.5}$	<sup>5</sup> GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$
$16.3 \pm 9.8^{+6.0}_{-2.0}$	<sup>6</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$
$13.3 \pm 3.3^{+4.0}_{-3.0}$	<sup>6</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$
$8.4 \pm 2.0 \pm 2.0$	<sup>6</sup> BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$
$20.9^{+5.4}_{-4.7} \pm 2.1_{-5.7}$	<sup>6</sup> BONDAR 12	BELL	$e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$
$19 \pm 7^{+11}_{-7}$	<sup>6</sup> BONDAR 12	BELL	$e^+e^- \rightarrow h_b(2P)\pi^+\pi^-$
<sup>4</sup> Average of the BONDAR 12 measurements in separate channels.			
<sup>5</sup> Correlated with the corresponding result from BONDAR 12.			
<sup>6</sup> Superseded by the average measurement of BONDAR 12.			

$X(10650)^+$  DECAY MODES

$X(10650)^-$  decay modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Upsilon(1S)\pi^+$	seen
$\Gamma_2 \quad \Upsilon(2S)\pi^+$	seen
$\Gamma_3 \quad \Upsilon(3S)\pi^+$	seen
$\Gamma_4 \quad h_b(1P)\pi^+$	seen
$\Gamma_5 \quad h_b(2P)\pi^+$	seen

$X(10650)^{\pm}$  BRANCHING RATIOS

$\Gamma(\Upsilon(1S)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$	
$\Gamma(\Upsilon(2S)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$	
$\Gamma(\Upsilon(3S)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	GARMASH 15	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$	
$\Gamma(h_b(1P)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow h_b(1P)\pi^+\pi^-$	
$\Gamma(h_b(2P)\pi^+)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	BONDAR 12	BELL	$e^+e^- \rightarrow h_b(2P)\pi^+\pi^-$	

$X(10650)^{\pm}$  REFERENCES

GARMASH 15	PR D91 072003	A. Garmash <i>et al.</i>	(BELLE Collab.)
BONDAR 12	PRL 108 122001	A. Bondar <i>et al.</i>	(BELLE Collab.)

# Meson Particle Listings

## $\Upsilon(10860)$

### $\Upsilon(10860)$

$$J^{PC} = 0^{-}(1^{-}-)$$

#### $\Upsilon(10860)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>10891.1 \pm 3.2^{+1.2}_{-2.0}</math></b>	<sup>1</sup> SANTEL	16 BELL	$e^{+}e^{-} \rightarrow \Upsilon(1S, 2S, 3S) \pi^{+} \pi^{-}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$10881.8^{+1.0}_{-1.1} \pm 1.2$	<sup>2,3</sup> SANTEL	16 BELL	$e^{+}e^{-} \rightarrow$ hadrons
$10879 \pm 3$	<sup>4,5</sup> CHEN	10 BELL	$e^{+}e^{-} \rightarrow$ hadrons
$10888.4^{+2.7}_{-2.6} \pm 1.2$	<sup>6</sup> CHEN	10 BELL	$e^{+}e^{-} \rightarrow \Upsilon(1S, 2S, 3S) \pi^{+} \pi^{-}$
$10876 \pm 2$	<sup>4</sup> AUBERT	09E BABR	$e^{+}e^{-} \rightarrow$ hadrons
$10869 \pm 2$	<sup>7</sup> AUBERT	09E BABR	$e^{+}e^{-} \rightarrow$ hadrons
$10868 \pm 6 \pm 5$	<sup>8</sup> BESSON	85 CLEO	$e^{+}e^{-} \rightarrow$ hadrons
$10845 \pm 20$	<sup>9</sup> LOVELOCK	85 CUSB	$e^{+}e^{-} \rightarrow$ hadrons

- <sup>1</sup> From a simultaneous fit to the  $\Upsilon(nS) \pi^{+} \pi^{-}$ ,  $n = 1, 2, 3$  cross sections at 25 energy points within  $\sqrt{s} = 10.6\text{--}11.05$  GeV to a pair of interfering Breit-Wigner amplitudes modified by phase space factors, with fourteen resonance parameters (a mass, width, and three amplitudes for each of  $\Upsilon(10860)$  and  $\Upsilon(11020)$ , a single universal relative phase, and three decoherence coefficients, one for each  $n$ ). Continuum contributions were measured (and therefore fixed) to be zero.
- <sup>2</sup> From a fit to the total hadronic cross sections measured at 60 energy points within  $\sqrt{s} = 10.82\text{--}11.05$  GeV to a pair of interfering Breit-Wigner amplitudes and two floating continuum amplitudes with  $1/\sqrt{s}$  dependence, one coherent with the resonances and one incoherent, with six resonance parameters (a mass, width, and an amplitude for each of  $\Upsilon(10860)$  and  $\Upsilon(11020)$ , one relative phase, and one decoherence coefficient).
- <sup>3</sup> Not including uncertain and potentially large systematic errors due to assumed continuum amplitude  $1/\sqrt{s}$  dependence and related interference contributions.
- <sup>4</sup> In a model where a flat non-resonant  $b\bar{b}$ -continuum is incoherently added to a second flat component interfering with two Breit-Wigner resonances. Systematic uncertainties not estimated.
- <sup>5</sup> The parameters of the  $\Upsilon(11020)$  are fixed to those in AUBERT 09E.
- <sup>6</sup> In a model where a flat nonresonant  $\Upsilon(1S, 2S, 3S) \pi^{+} \pi^{-}$  continuum interferes with a single Breit-Wigner resonance.
- <sup>7</sup> In a model where a non-resonant  $b\bar{b}$ -continuum represented by a threshold function at  $\sqrt{s}=2m_B$  is incoherently added to a flat component interfering with two Breit-Wigner resonances. Not independent of other AUBERT 09E results. Systematic uncertainties not estimated.
- <sup>8</sup> Assuming four Gaussians with radiative tails and a single step in  $R$ .
- <sup>9</sup> In a coupled-channel model with three resonances and a smooth step in  $R$ .

#### $\Upsilon(10860)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>53.7^{+7.1+1.3}_{-5.6-5.4}</math></b>	<sup>10</sup> SANTEL	16 BELL	$e^{+}e^{-} \rightarrow \Upsilon(1S, 2S, 3S) \pi^{+} \pi^{-}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$48.5^{+1.9+2.0}_{-1.8-2.8}$	<sup>11,12</sup> SANTEL	16 BELL	$e^{+}e^{-} \rightarrow$ hadrons
$46^{+9}_{-7}$	<sup>13,14</sup> CHEN	10 BELL	$e^{+}e^{-} \rightarrow$ hadrons
$30.7^{+8.3}_{-7.0} \pm 3.1$	<sup>15</sup> CHEN	10 BELL	$e^{+}e^{-} \rightarrow \Upsilon(1S, 2S, 3S) \pi^{+} \pi^{-}$
$43 \pm 4$	<sup>13</sup> AUBERT	09E BABR	$e^{+}e^{-} \rightarrow$ hadrons
$74 \pm 4$	<sup>16</sup> AUBERT	09E BABR	$e^{+}e^{-} \rightarrow$ hadrons
$112 \pm 17 \pm 23$	<sup>17</sup> BESSON	85 CLEO	$e^{+}e^{-} \rightarrow$ hadrons
$110 \pm 15$	<sup>18</sup> LOVELOCK	85 CUSB	$e^{+}e^{-} \rightarrow$ hadrons

- <sup>10</sup> From a simultaneous fit to the  $\Upsilon(nS) \pi^{+} \pi^{-}$ ,  $n = 1, 2, 3$  cross sections at 25 energy points within  $\sqrt{s} = 10.6\text{--}11.05$  GeV to a pair of interfering Breit-Wigner amplitudes modified by phase space factors, with fourteen resonance parameters (a mass, width, and three amplitudes for each of  $\Upsilon(10860)$  and  $\Upsilon(11020)$ , a single universal relative phase, and three decoherence coefficients, one for each  $n$ ). Continuum contributions were measured (and therefore fixed) to be zero.
- <sup>11</sup> From a fit to the total hadronic cross sections measured at 60 energy points within  $\sqrt{s} = 10.82\text{--}11.05$  GeV to a pair of interfering Breit-Wigner amplitudes and two floating continuum amplitudes with  $1/\sqrt{s}$  dependence, one coherent with the resonances and one incoherent, with six resonance parameters (a mass, width, and an amplitude for each of  $\Upsilon(10860)$  and  $\Upsilon(11020)$ , one relative phase, and one decoherence coefficient).
- <sup>12</sup> Not including uncertain and potentially large systematic errors due to assumed continuum amplitude  $1/\sqrt{s}$  dependence and related interference contributions.
- <sup>13</sup> In a model where a flat non-resonant  $b\bar{b}$ -continuum is incoherently added to a second flat component interfering with two Breit-Wigner resonances. Systematic uncertainties not estimated.
- <sup>14</sup> The parameters of the  $\Upsilon(11020)$  are fixed to those in AUBERT 09E.
- <sup>15</sup> In a model where a flat nonresonant  $\Upsilon(1S, 2S, 3S) \pi^{+} \pi^{-}$  continuum interferes with a single Breit-Wigner resonance.
- <sup>16</sup> In a model where a non-resonant  $b\bar{b}$ -continuum represented by a threshold function at  $\sqrt{s}=2m_B$  is incoherently added to a flat component interfering with two Breit-Wigner resonances. Not independent of other AUBERT 09E results. Systematic uncertainties not estimated.
- <sup>17</sup> Assuming four Gaussians with radiative tails and a single step in  $R$ .
- <sup>18</sup> In a coupled-channel model with three resonances and a smooth step in  $R$ .

#### $\Upsilon(10860)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $B\bar{B}X$	( $76.2^{+2.7}_{-4.0}$ ) %	
$\Gamma_2$ $B\bar{B}$	( $5.5 \pm 1.0$ ) %	
$\Gamma_3$ $B\bar{B}^{*} + \text{c.c.}$	( $13.7 \pm 1.6$ ) %	
$\Gamma_4$ $B^{*}\bar{B}^{*}$	( $38.1 \pm 3.4$ ) %	
$\Gamma_5$ $B\bar{B}^{(*)}\pi$	< 19.7 %	90%
$\Gamma_6$ $B\bar{B}\pi$	( $0.0 \pm 1.2$ ) %	
$\Gamma_7$ $B^{*}\bar{B}\pi + B\bar{B}^{*}\pi$	( $7.3 \pm 2.3$ ) %	
$\Gamma_8$ $B^{*}\bar{B}^{*}\pi$	( $1.0 \pm 1.4$ ) %	
$\Gamma_9$ $B\bar{B}\pi\pi$	< 8.9 %	90%
$\Gamma_{10}$ $B_s^{(*)}\bar{B}_s^{(*)}$	( $20.1 \pm 3.1$ ) %	
$\Gamma_{11}$ $B_s\bar{B}_s$	( $5 \pm 5$ ) $\times 10^{-3}$	
$\Gamma_{12}$ $B_s\bar{B}_s^{*} + \text{c.c.}$	( $1.35 \pm 0.32$ ) %	
$\Gamma_{13}$ $B_s^{*}\bar{B}_s^{*}$	( $17.6 \pm 2.7$ ) %	
$\Gamma_{14}$ no open-bottom	( $3.8^{+5.0}_{-0.5}$ ) %	
$\Gamma_{15}$ $e^{+}e^{-}$	( $5.7 \pm 1.5$ ) $\times 10^{-6}$	
$\Gamma_{16}$ $K^{*}(892)^0\bar{K}^0$	< 1.0 $\times 10^{-5}$	90%
$\Gamma_{17}$ $\Upsilon(1S) \pi^{+} \pi^{-}$	( $5.3 \pm 0.6$ ) $\times 10^{-3}$	
$\Gamma_{18}$ $\Upsilon(2S) \pi^{+} \pi^{-}$	( $7.8 \pm 1.3$ ) $\times 10^{-3}$	
$\Gamma_{19}$ $\Upsilon(3S) \pi^{+} \pi^{-}$	( $4.8^{+1.9}_{-1.7}$ ) $\times 10^{-3}$	
$\Gamma_{20}$ $\Upsilon(1S) K^{+} K^{-}$	( $6.1 \pm 1.8$ ) $\times 10^{-4}$	
$\Gamma_{21}$ $h_b(1P) \pi^{+} \pi^{-}$	( $3.5^{+1.0}_{-1.3}$ ) $\times 10^{-3}$	
$\Gamma_{22}$ $h_b(2P) \pi^{+} \pi^{-}$	( $6.0^{+2.1}_{-1.8}$ ) $\times 10^{-3}$	
$\Gamma_{23}$ $\chi_{b0}(1P) \pi^{+} \pi^{-} \pi^0$	< 6.3 $\times 10^{-3}$	90%
$\Gamma_{24}$ $\chi_{b0}(1P) \omega$	< 3.9 $\times 10^{-3}$	90%
$\Gamma_{25}$ $\chi_{b0}(1P) (\pi^{+} \pi^{-} \pi^0)_{\text{non-}\omega}$	< 4.8 $\times 10^{-3}$	90%
$\Gamma_{26}$ $\chi_{b1}(1P) \pi^{+} \pi^{-} \pi^0$	( $1.85 \pm 0.33$ ) $\times 10^{-3}$	
$\Gamma_{27}$ $\chi_{b1}(1P) \omega$	( $1.57 \pm 0.30$ ) $\times 10^{-3}$	
$\Gamma_{28}$ $\chi_{b1}(1P) (\pi^{+} \pi^{-} \pi^0)_{\text{non-}\omega}$	( $5.2 \pm 1.9$ ) $\times 10^{-4}$	
$\Gamma_{29}$ $\chi_{b2}(1P) \pi^{+} \pi^{-} \pi^0$	( $1.17 \pm 0.30$ ) $\times 10^{-3}$	
$\Gamma_{30}$ $\chi_{b2}(1P) \omega$	( $6.0 \pm 2.7$ ) $\times 10^{-4}$	
$\Gamma_{31}$ $\chi_{b2}(1P) (\pi^{+} \pi^{-} \pi^0)_{\text{non-}\omega}$	( $6 \pm 4$ ) $\times 10^{-4}$	
$\Gamma_{32}$ $\gamma \chi_b \rightarrow \gamma \Upsilon(1S) \omega$	< 3.8 $\times 10^{-5}$	90%

#### Inclusive Decays.

These decay modes are submodes of one or more of the decay modes above.

$\Gamma_{33}$ $\phi$ anything	( $13.8^{+2.4}_{-1.7}$ ) %
$\Gamma_{34}$ $D^0$ anything + c.c.	( $108 \pm 8$ ) %
$\Gamma_{35}$ $D_s$ anything + c.c.	( $46 \pm 6$ ) %
$\Gamma_{36}$ $J/\psi$ anything	( $2.06 \pm 0.21$ ) %
$\Gamma_{37}$ $B^0$ anything + c.c.	( $77 \pm 8$ ) %
$\Gamma_{38}$ $B^{+}$ anything + c.c.	( $72 \pm 6$ ) %

#### $\Upsilon(10860)$ PARTIAL WIDTHS

$\Gamma(e^{+}e^{-})$	VALUE (keV)	DOCUMENT ID	TECN	COMMENT	$\Gamma_{15}$
<b><math>0.31 \pm 0.07</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.			
	$0.22 \pm 0.05 \pm 0.07$	BESSON	85 CLEO	$e^{+}e^{-} \rightarrow$ hadrons	
	$0.365 \pm 0.070$	LOVELOCK	85 CUSB	$e^{+}e^{-} \rightarrow$ hadrons	

#### $\Upsilon(10860)$ BRANCHING RATIOS

“OUR EVALUATION” is obtained based on averages of rescaled data listed below. The averages and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>.

$\Gamma(B\bar{B}X)/\Gamma_{\text{total}}$	VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b><math>0.762^{+0.027}_{-0.043}</math> OUR EVALUATION</b>						
<b><math>0.71 \pm 0.06</math> OUR AVERAGE</b>						
	$0.737 \pm 0.032 \pm 0.051$	1063	<sup>19</sup> DRUTSKOY	10 BELL	$\Upsilon(5S) \rightarrow B^{+} X, B^0 X$	
	$0.589 \pm 0.100 \pm 0.092$		<sup>20</sup> HUANG	07 CLEO	$\Upsilon(5S) \rightarrow$ hadrons	

$\Gamma(B\bar{B})/\Gamma_{\text{total}}$	VALUE (units $10^{-2}$ )	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
<b><math>5.5^{+1.0}_{-0.9} \pm 0.4</math></b>			<sup>21</sup> DRUTSKOY	10 BELL	$\Upsilon(5S) \rightarrow B^{+} X, B^0 X$	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<13.8		90	<sup>20</sup> HUANG	07 CLEO	$\Upsilon(5S) \rightarrow$ hadrons	

See key on page 601

Meson Particle Listings  
 $\Upsilon(10860)$ 

$\Gamma(B\bar{B})/\Gamma(B\bar{B}X)$		$\Gamma_2/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.22</b>	90	AQUINES	06 CLE3 $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B\bar{B}^* + \text{c.c.})/\Gamma_{\text{total}}$		$\Gamma_3/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.137±0.016 OUR AVERAGE</b>			
0.137±0.013±0.011	21 DRUTSKOY	10 BELL	$\Upsilon(5S) \rightarrow B^+X, B^0X$
0.143±0.053±0.027	20 HUANG	07 CLEO	$\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B\bar{B}^* + \text{c.c.})/\Gamma(B\bar{B}X)$		$\Gamma_3/\Gamma_1$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.24±0.09±0.03</b>	10	AQUINES	06 CLE3 $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B^*\bar{B}^*)/\Gamma_{\text{total}}$		$\Gamma_4/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.381±0.034 OUR AVERAGE</b>			
0.375 $^{+0.021}_{-0.019}$ ±0.030	21 DRUTSKOY	10 BELL	$\Upsilon(5S) \rightarrow B^+X, B^0X$
0.436±0.083±0.072	20 HUANG	07 CLEO	$\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B^*\bar{B}^*)/\Gamma(B\bar{B}X)$		$\Gamma_4/\Gamma_1$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.74±0.15±0.08</b>	31	AQUINES	06 CLE3 $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B\bar{B}^{(*)}\pi)/\Gamma_{\text{total}}$		$\Gamma_5/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.197</b>	90	20 HUANG	07 CLEO $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B\bar{B}^{(*)}\pi)/\Gamma(B\bar{B}X)$		$\Gamma_5/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.32</b>	90	AQUINES	06 CLE3 $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B\bar{B}\pi)/\Gamma_{\text{total}}$		$\Gamma_6/\Gamma$	
VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.0±1.2±0.3</b>	0	21 DRUTSKOY	10 BELL $\Upsilon(5S) \rightarrow B^{+0}\pi^-X$

$[\Gamma(B^*\bar{B}\pi) + \Gamma(B\bar{B}^*\pi)]/\Gamma_{\text{total}}$		$\Gamma_7/\Gamma$	
VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>7.3<math>^{+2.3}_{-2.1}</math>±0.8</b>	38	21 DRUTSKOY	10 BELL $\Upsilon(5S) \rightarrow B^{+0}\pi^-X$

$\Gamma(B^*\bar{B}^*\pi)/\Gamma_{\text{total}}$		$\Gamma_8/\Gamma$	
VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>1.0<math>^{+1.4}_{-1.3}</math>±0.4</b>	5	21 DRUTSKOY	10 BELL $\Upsilon(5S) \rightarrow B^{+0}\pi^-X$

$\Gamma(B\bar{B}\pi\pi)/\Gamma_{\text{total}}$		$\Gamma_9/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.089</b>	90	20 HUANG	07 CLEO $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B\bar{B}\pi\pi)/\Gamma(B\bar{B}X)$		$\Gamma_9/\Gamma_1$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.14</b>	90	AQUINES	06 CLE3 $\Upsilon(5S) \rightarrow \text{hadrons}$

$\Gamma(B_s^{(*)}\bar{B}_s^{(*)})/\Gamma_{\text{total}}$		$\Gamma_{10}/\Gamma = \Gamma_{11} + \Gamma_{12} + \Gamma_{13}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.201<math>^{+0.030}_{-0.031}</math> OUR EVALUATION</b>			

<b>0.189<math>^{+0.027}_{-0.021}</math> OUR AVERAGE</b>			
0.172±0.030	22 ESEN	13 BELL	$\Upsilon(5S) \rightarrow D^0X, D_sX$
0.21 $^{+0.06}_{-0.03}$	23 HUANG	07 CLEO	$\Upsilon(5S) \rightarrow D_sX$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.180±0.013±0.032	24 DRUTSKOY	07 BELL	$\Upsilon(5S) \rightarrow D^0X, D_sX$
0.160±0.026±0.058	25 ARTUSO	05B CLEO	$e^+e^- \rightarrow D_XX$

$\Gamma(B_s^{(*)}\bar{B}_s^{(*)})/\Gamma(B\bar{B}X)$		$\Gamma_{10}/\Gamma_1$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.264<math>^{+0.052}_{-0.045}</math> OUR EVALUATION</b>			

$\Gamma(B_s^*\bar{B}_s^*)/\Gamma(B_s^{(*)}\bar{B}_s^{(*)})$		$\Gamma_{13}/\Gamma_{10} = \Gamma_{13}/(\Gamma_{11} + \Gamma_{12} + \Gamma_{13})$	
VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>87.8±1.5 OUR AVERAGE</b>			
87.0±1.7	26,27 ESEN	13 BELL	$B_s^0 \rightarrow D_s^- \pi^+$
90.5±3.2±0.1	227 27,28 LI	12 BELL	$B_s^0 \rightarrow J/\psi \eta^{(\prime)}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
90.1 $^{+3.8}_{-4.0}$ ±0.2	29 LOUVOT	09 BELL	10.86 $e^+e^- \rightarrow B_s^{(*)}\bar{B}_s^{(*)}$
93 $^{+7}_{-9}$ ±1	29 DRUTSKOY	07A BELL	Superseded by LOUVOT 09

$\Gamma(B_s\bar{B}_s)/\Gamma(B_s^{(*)}\bar{B}_s^{(*)})$		$\Gamma_{11}/\Gamma_{10} = \Gamma_{11}/(\Gamma_{11} + \Gamma_{12} + \Gamma_{13})$	
VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>2.6<math>^{+2.6}_{-2.5}</math></b>	LOUVOT	09 BELL	10.86 $e^+e^- \rightarrow B_s^{(*)}\bar{B}_s^{(*)}$

$\Gamma(B_s\bar{B}_s)/\Gamma(B_s^*\bar{B}_s^*)$		$\Gamma_{11}/\Gamma_{13}$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.16</b>	90	BONVICINI	06 CLE3 $e^+e^-$

$\Gamma(B_s\bar{B}_s^* + \text{c.c.})/\Gamma(B_s^{(*)}\bar{B}_s^{(*)})$		$\Gamma_{12}/\Gamma_{10} = \Gamma_{12}/(\Gamma_{11} + \Gamma_{12} + \Gamma_{13})$	
VALUE (units 10 <sup>-2</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>6.7±1.2 OUR AVERAGE</b>			
7.3±1.4	26,27 ESEN	13 BELL	$B_s^0 \rightarrow D_s^- \pi^+$
4.9±2.5±0.0	227 27,28 LI	12 BELL	$B_s^0 \rightarrow J/\psi \eta^{(\prime)}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
7.3 $^{+3.3}_{-3.0}$ ±0.1	LOUVOT	09 BELL	10.86 $e^+e^- \rightarrow B_s^{(*)}\bar{B}_s^{(*)}$

$\Gamma(B_s\bar{B}_s^* + \text{c.c.})/\Gamma(B_s^*\bar{B}_s^*)$		$\Gamma_{12}/\Gamma_{13}$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.16</b>	90	BONVICINI	06 CLE3 $e^+e^-$

$\Gamma(\text{no open-bottom})/\Gamma_{\text{total}}$		$\Gamma_{14}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.038<math>^{+0.051}_{-0.005}</math> OUR EVALUATION</b>			

$\Gamma(K^*(892)^0\bar{K}^0)/\Gamma_{\text{total}}$		$\Gamma_{16}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;1.0 × 10<sup>-5</sup></b>	90	SHEN	13A BELL $e^+e^- \rightarrow K^*(892)^0\bar{K}^0$

$\Gamma(\Upsilon(1S)\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{17}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>5.3±0.3±0.5</b>	325	30 CHEN	08 BELL 10.87 $e^+e^- \rightarrow \Upsilon(1S)\pi^+\pi^-$

$\Gamma(\Upsilon(2S)\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{18}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>7.8±0.6±1.1</b>	186	30 CHEN	08 BELL 10.87 $e^+e^- \rightarrow \Upsilon(2S)\pi^+\pi^-$

$\Gamma(\Upsilon(3S)\pi^+\pi^-)/\Gamma_{\text{total}}$		$\Gamma_{19}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>4.8<math>^{+1.8}_{-1.5}</math>±0.7</b>	10	30 CHEN	08 BELL 10.87 $e^+e^- \rightarrow \Upsilon(3S)\pi^+\pi^-$

$\Gamma(\Upsilon(1S)K^+K^-)/\Gamma_{\text{total}}$		$\Gamma_{20}/\Gamma$	
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>6.1<math>^{+1.6}_{-1.4}</math>±1.0</b>	20	30 CHEN	08 BELL 10.87 $e^+e^- \rightarrow \Upsilon(1S)K^+K^-$

$\Gamma(h_b(1P)\pi^+\pi^-)/\Gamma(\Upsilon(2S)\pi^+\pi^-)$		$\Gamma_{21}/\Gamma_{18}$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.45±0.08<math>^{+0.07}_{-0.12}</math></b>	ADACHI	12 BELL	10.86 $e^+e^- \rightarrow \text{hadrons}$

$\Gamma(h_b(2P)\pi^+\pi^-)/\Gamma(\Upsilon(2S)\pi^+\pi^-)$		$\Gamma_{22}/\Gamma_{18}$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.77±0.08<math>^{+0.22}_{-0.17}</math></b>	ADACHI	12 BELL	10.86 $e^+e^- \rightarrow \text{hadrons}$

$\Gamma(\chi_{b0}(1P)\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{23}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;6.3 × 10<sup>-3</sup></b>	90	31 HE	14 BELL $\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$

$\Gamma(\chi_{b0}(1P)\omega)/\Gamma_{\text{total}}$		$\Gamma_{24}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;3.9 × 10<sup>-3</sup></b>	90	31 HE	14 BELL $\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$

$\Gamma(\chi_{b0}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega})/\Gamma_{\text{total}}$		$\Gamma_{25}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;4.8 × 10<sup>-3</sup></b>	90	31 HE	14 BELL $\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$

$\Gamma(\chi_{b1}(1P)\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$		$\Gamma_{26}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>1.85±0.23±0.23</b>	80	31 HE	14 BELL $\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$

$\Gamma(\chi_{b1}(1P)\omega)/\Gamma_{\text{total}}$		$\Gamma_{27}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>1.57±0.22±0.21</b>	60	31 HE	14 BELL $\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$

$\Gamma(\chi_{b1}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega})/\Gamma_{\text{total}}$		$\Gamma_{28}/\Gamma$	
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.52±0.15±0.11</b>	24	31 HE	14 BELL $\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$

# Meson Particle Listings

## $\Upsilon(10860)$ , $\Upsilon(11020)$

$\Gamma(\chi_{b2}(1P)\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
$1.17\pm0.27\pm0.14$	29	31 HE	14	BELL	$\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$	
$\Gamma(\chi_{b2}(1P)\omega)/\Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
$0.60\pm0.23\pm0.15$	13	31 HE	14	BELL	$\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$	
$\Gamma(\chi_{b2}(1P)\omega)/\Gamma(\chi_{b1}(1P)\omega)$					$\Gamma_{30}/\Gamma_{27}$	
VALUE		DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.38\pm0.16\pm0.09$	32	HE	14	BELL	$\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$	
$\Gamma(\chi_{b2}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega})/\Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
$0.61\pm0.22\pm0.28$	16	31 HE	14	BELL	$\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$	
$\Gamma(\chi_{b2}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega})/\Gamma(\chi_{b1}(1P)(\pi^+\pi^-\pi^0)_{\text{non-}\omega})$					$\Gamma_{31}/\Gamma_{28}$	
VALUE		DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$1.20\pm0.55\pm0.65$	32	HE	14	BELL	$\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$	
$\Gamma(\gamma X_b \rightarrow \gamma\Upsilon(1S)\omega)/\Gamma_{\text{total}}$					$\Gamma_{32}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		
$<3.8 \times 10^{-5}$	90	33 HE	14	BELL	$\Upsilon(5S) \rightarrow \pi^+\pi^-\pi^0\gamma\Upsilon(1S)$	
$\Gamma(\phi \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_{33}/\Gamma$	
VALUE		DOCUMENT ID	TECN	COMMENT		
$0.138\pm0.007\pm0.023_{-0.015}$		HUANG	07	CLEO	$\Upsilon(5S) \rightarrow \phi X$	
$\Gamma(D^0 \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{34}/\Gamma$	
VALUE		DOCUMENT ID	TECN	COMMENT		
$1.076\pm0.040\pm0.068$		DRUTSKOY	07	BELL	$\Upsilon(5S) \rightarrow D^0 X$	
$\Gamma(D_s \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{35}/\Gamma$	
VALUE		DOCUMENT ID	TECN	COMMENT		
$0.46\pm0.06$ OUR AVERAGE						
$0.472\pm0.024\pm0.072$	24	DRUTSKOY	07	BELL	$\Upsilon(5S) \rightarrow D_s X$	
$0.44\pm0.09\pm0.04$	34	ARTUSO	05B	CLE3	$e^+e^- \rightarrow D_X X$	
$\Gamma(J/\psi \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_{36}/\Gamma$	
VALUE (units $10^{-2}$ )		DOCUMENT ID	TECN	COMMENT		
$2.060\pm0.160\pm0.134$		DRUTSKOY	07	BELL	$\Upsilon(5S) \rightarrow J/\psi X$	
$\Gamma(B^0 \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{37}/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
$0.770\pm0.058_{-0.056}\pm0.061$	352	DRUTSKOY	10	BELL	$\Upsilon(5S) \rightarrow B^0 X$	
$\Gamma(B^+ \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}$					$\Gamma_{38}/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
$0.721\pm0.039_{-0.038}\pm0.050$	711	DRUTSKOY	10	BELL	$\Upsilon(5S) \rightarrow B^+ X$	

<sup>19</sup>Not independent of DRUTSKOY 10 values for  $\Upsilon(5S) \rightarrow B^{\pm,0} \text{ anything}$ .

<sup>20</sup>Using measurements or limits from AQUINES 06.

<sup>21</sup>Assuming isospin conservation.

<sup>22</sup>Supersedes DRUTSKOY 07.

<sup>23</sup>Supersedes ARTUSO 05B. Combining inclusive  $\phi$ ,  $D_s$ , and  $B$  measurements. Using  $B(D_s^+ \rightarrow \phi\pi^+) = 4.4 \pm 0.6\%$  from PDG 06.

<sup>24</sup>Using  $B(D_s^+ \rightarrow \phi\pi^+) = (4.4 \pm 0.6)\%$  from PDG 06.

<sup>25</sup>Uses a model-dependent estimate  $B(B_S \rightarrow D_s X) = (92 \pm 11)\%$ .

<sup>26</sup>Supersedes LOUVOT 09.

<sup>27</sup>With  $N(B_S^{(*)}\overline{B}_S^{(*)}) = (7.11 \pm 1.30) \times 10^6$ .

<sup>28</sup>The ratios  $N(B_S^{(*)}\overline{B}_S^{(*)})/N(B_S^{(*)}\overline{B}_S^{(*)})$  and  $N(B_S^{(*)}\overline{B}_S^0)/N(B_S^{(*)}\overline{B}_S^{(*)})$  are measured with a correlation coefficient of  $-0.72$ .

<sup>29</sup>From a measurement of  $\sigma(e^+e^- \rightarrow B_S^* \overline{B}_S^*)/\sigma(e^+e^- \rightarrow B_S^{(*)} \overline{B}_S^{(*)})$  at  $\sqrt{s} = 10.86$  GeV.

<sup>30</sup>Assuming that the observed events are solely due to the  $\Upsilon(5S)$  resonance.

<sup>31</sup>Assuming that all the  $b\overline{b}$  events are from  $\Upsilon(5S)$  resonance decays and using  $\sigma(e^+e^- \rightarrow b\overline{b}) = 0.340 \pm 0.016$  nb from ESEN 13. Correlated with other results from HE 14.

<sup>32</sup>Accounting for correlated systematics.

<sup>33</sup>Assuming that all the  $b\overline{b}$  events are from  $\Upsilon(5S)$  resonance decays and using  $\sigma(e^+e^- \rightarrow b\overline{b}) = 0.340 \pm 0.016$  nb from ESEN 13. Correlated with other results from HE 14. For a state  $X_b$  with mass between  $10.55$  GeV/ $c^2$  and  $10.65$  GeV/ $c^2$ , the obtained 90% upper limit as a function of  $m_{X_b}$  varies from  $2.6 \times 10^{-5}$  to  $3.8 \times 10^{-5}$ .

<sup>34</sup>ARTUSO 05B reports  $[\Gamma(\Upsilon(10860) \rightarrow D_s \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)] = 0.0198 \pm 0.0019 \pm 0.0038$  which we divide by our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>19</sup> Not independent of DRUTSKOY 10 values for  $\Upsilon(5S) \rightarrow B^{\pm,0} \text{ anything}$ .  
<sup>20</sup> Using measurements or limits from AQUINES 06.  
<sup>21</sup> Assuming isospin conservation.  
<sup>22</sup> Supersedes DRUTSKOY 07.  
<sup>23</sup> Supersedes ARTUSO 05B. Combining inclusive  $\phi$ ,  $D_s$ , and  $B$  measurements. Using  $B(D_s^+ \rightarrow \phi\pi^+) = 4.4 \pm 0.6\%$  from PDG 06.  
<sup>24</sup> Using  $B(D_s^+ \rightarrow \phi\pi^+) = (4.4 \pm 0.6)\%$  from PDG 06.  
<sup>25</sup> Uses a model-dependent estimate  $B(B_s \rightarrow D_s X) = (92 \pm 11)\%$ .  
<sup>26</sup> Supersedes LOUVOT 09.  
<sup>27</sup> With  $N(B_s^{(*)}\bar{B}_s^{(*)}) = (7.11 \pm 1.30) \times 10^6$ .  
<sup>28</sup> The ratios  $N(B_s^{(*)}\bar{B}_s^{(*)})/N(B_s^{(*)}\bar{B}_s^{(*)})$  and  $N(B_s^{(*)}\bar{B}_s^{(*)})/N(B_s^{(*)}\bar{B}_s^{(*)})$  are measured with a correlation coefficient of  $-0.72$ .  
<sup>29</sup> From a measurement of  $\sigma(e^+e^- \rightarrow B_s^*\bar{B}_s^*)/\sigma(e^+e^- \rightarrow B_s^{(*)}\bar{B}_s^{(*)})$  at  $\sqrt{s} = 10.86$  GeV.  
<sup>30</sup> Assuming that the observed events are solely due to the  $\Upsilon(5S)$  resonance.  
<sup>31</sup> Assuming that all the  $b\bar{b}$  events are from  $\Upsilon(5S)$  resonance decays and using  $\sigma(e^+e^- \rightarrow b\bar{b}) = 0.340 \pm 0.016$  nb from ESEN 13. Correlated with other results from HE 14.  
<sup>32</sup> Accounting for correlated systematics.  
<sup>33</sup> Assuming that all the  $b\bar{b}$  events are from  $\Upsilon(5S)$  resonance decays and using  $\sigma(e^+e^- \rightarrow b\bar{b}) = 0.340 \pm 0.016$  nb from ESEN 13. Correlated with other results from HE 14. For a state  $X_b$  with mass between 10.55 GeV/c<sup>2</sup> and 10.65 GeV/c<sup>2</sup>, the obtained 90% upper limit as a function of  $m_{X_b}$  varies from  $2.6 \times 10^{-5}$  to  $3.8 \times 10^{-5}$ .  
<sup>34</sup> ARTUSO 05B reports  $[\Gamma(\Upsilon(10860) \rightarrow D_s \text{ anything} + \text{c.c.})/\Gamma_{\text{total}}] \times [B(D_s^+ \rightarrow \phi\pi^+)] = 0.0198 \pm 0.0019 \pm 0.0038$  which we divide by our best value  $B(D_s^+ \rightarrow \phi\pi^+) = (4.5 \pm 0.4) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Upsilon(10860)$ REFERENCES									
SANTEL	16	PR D93 011101	D. Santel <i>et al.</i>	(BELLE Collab.)					
HE	14	PRL 113 142001	X.H. He <i>et al.</i>	(BELLE Collab.)					
ESEN	13	PR D87 031101	S. ESEN <i>et al.</i>	(BELLE Collab.)					
SHEN	13A	PR D88 052019	C.P. Shen <i>et al.</i>	(BELLE Collab.)					
ADACHI	12	PRL 108 032001	I. Adachi <i>et al.</i>	(BELLE Collab.)					
LI	12	PRL 108 181808	J. Li <i>et al.</i>	(BELLE Collab.)					
CHEN	10	PR D82 091106	K.-F. Chen <i>et al.</i>	(BELLE Collab.)					
DRUTSKOY	10	PR D81 112003	A. Drutskoy <i>et al.</i>	(BELLE Collab.)					
AUBERT	09E	PRL 102 012001	B. Aubert <i>et al.</i>	(BABAR Collab.)					
LOUVOT	09	PRL 102 021801	R. Louvot <i>et al.</i>	(BELLE Collab.)					
CHEN	08	PRL 100 112001	K.-F. Chen <i>et al.</i>	(BELLE Collab.)					
DRUTSKOY	07	PR 98 052001	A. Drutskoy <i>et al.</i>	(BELLE Collab.)					
DRUTSKOY	07A	PR D76 012002	A. Drutskoy <i>et al.</i>	(BELLE Collab.)					
HUANG	07	PR D75 012002	G.S. Huang <i>et al.</i>	(CLEO Collab.)					
AQUINES	06	PRL 96 152001	O. Aquines <i>et al.</i>	(CLEO Collab.)					
BONVICINI	06	PRL 96 022002	G. Bonvicini <i>et al.</i>	(CLEO Collab.)					
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)					
ARTUSO	05B	PRL 95 261801	M. Artuso <i>et al.</i>	(CLEO Collab.)					
BESSION	85	PRL 54 381	D. Besson <i>et al.</i>	(CLEO Collab.)					
LOVELOCK	85	PRL 54 377	D.M.J. Lovelock <i>et al.</i>	(CUSB Collab.)					

$\Upsilon(11020)$

$\Gamma(G(J^PC)) = 0^-(1^--)$

$\Upsilon(11020)$ MASS					
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT		
$10987.5 \pm \begin{smallmatrix} 6.4+9.1 \\ 2.5-2.3 \end{smallmatrix}$	<sup>1</sup> SANTEL	16	BELL	$e^+e^- \rightarrow \Upsilon(1S, 2S, 3S)\pi^+\pi^-$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$11003.0 \pm \begin{smallmatrix} 1.1+0.9 \\ -1.0 \end{smallmatrix}$	<sup>2,3</sup> SANTEL	16	BELL	$e^+e^- \rightarrow \text{hadrons}$	
$10996 \pm 2$	<sup>4</sup> AUBERT	09E	BABR	$e^+e^- \rightarrow \text{hadrons}$	
$11019 \pm 5 \pm 7$	BESSION	85	CLEO	$e^+e^- \rightarrow \text{hadrons}$	
$11020 \pm 30$	LOVELOCK	85	CUSB	$e^+e^- \rightarrow \text{hadrons}$	
<sup>1</sup> From a simultaneous fit to the $\Upsilon(nS)\pi^+\pi^-$ , $n = 1, 2, 3$ cross sections at 25 energy points within $\sqrt{s} = 10.6\text{--}11.05$ GeV to a pair of interfering Breit-Wigner amplitudes modified by phase space factors, with fourteen resonance parameters (a mass, width, and three amplitudes for each of $\Upsilon(10860)$ and $\Upsilon(11020)$ , a single universal relative phase, and three decoherence coefficients, one for each $n$ ). Continuum contributions were measured (and therefore fixed) to be zero.					
<sup>2</sup> From a fit to the total hadronic cross sections measured at 60 energy points within $\sqrt{s} = 10.82\text{--}11.05$ GeV to a pair of interfering Breit-Wigner amplitudes and two floating continuum amplitudes with $1/\sqrt{s}$ dependence, one coherent with the resonances and one incoherent, with six resonance parameters (a mass, width, and an amplitude for each of $\Upsilon(10860)$ and $\Upsilon(11020)$ , one relative phase, and one decoherence coefficient).					
<sup>3</sup> Not including uncertain and potentially large systematic errors due to assumed continuum amplitude $1/\sqrt{s}$ dependence and related interference contributions.					
<sup>4</sup> In a model where a flat non-resonant $b\bar{b}$ -continuum is incoherently added to a second flat component interfering with two Breit-Wigner resonances. Systematic uncertainties not estimated.					

$\Upsilon(11020)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
$61 \pm 9 \pm 20$	<sup>5</sup> SANTEL	16	BELL	$e^+e^- \rightarrow \Upsilon(1S, 2S, 3S)\pi^+\pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$39.3 \pm 1.7 \pm 1.3$ $-1.6 - 2.4$	<sup>6,7</sup> SANTEL	16	BELL	$e^+e^- \rightarrow$ hadrons
$37 \pm 3$	<sup>8</sup> AUBERT	09E	BABR	$e^+e^- \rightarrow$ hadrons
$61 \pm 13 \pm 22$	BESSION	85	CLEO	$e^+e^- \rightarrow$ hadrons
$90 \pm 20$	LOVELOCK	85	CUSB	$e^+e^- \rightarrow$ hadrons

$\Upsilon(11020)$ PARTIAL WIDTHS				$\Upsilon(11020)$ REFERENCES			
$\Gamma(e^+e^-)$							
VALUE (keV)	DOCUMENT ID	TECN	COMMENT				
<b>0.130±0.030 OUR AVERAGE</b>				SANTEL	16	PR D93 011101	D. Santel <i>et al.</i>
0.095±0.03 ±0.035	BESSON	85	CLEO $e^+e^- \rightarrow$ hadrons	AUBERT	09E	PRL 102 012001	B. Aubert <i>et al.</i>
0.156±0.040	LOVELOCK	85	CUSB $e^+e^- \rightarrow$ hadrons	BESSON	85	PRL 54 381	D. Besson <i>et al.</i>
				LOVELOCK	85	PRL 54 377	D.M.J. Lovelock <i>et al.</i>

# Meson Particle Listings

## Non- $q\bar{q}$ Candidates

### NON- $q\bar{q}$ CANDIDATES

#### NON- $q\bar{q}$ MESONS

Revised August 2015 by C. Amsler (University of Bern and Stefan Meyer Institute for Subatomic Physics, Vienna) and C. Hanhart (Forschungszentrum Jülich).

The constituent quark model describes the observed meson spectrum as bound  $q\bar{q}$  states grouped into SU(N) flavor multiplets (see our review on the ‘Quark Model’ in this issue of the *Review*). However, the self-coupling of gluons in QCD suggests that additional mesons made of bound gluons (glueballs), or  $q\bar{q}$ -pairs with an excited gluon (hybrids), may exist. Multiquark color singlet states such as  $qq\bar{q}\bar{q}$  (tetraquarks as compact diquark-antidiquark systems and ‘molecular’ bound states of two mesons) or  $qqq\bar{q}\bar{q}\bar{q}$  (six-quark and ‘baryonium’ bound states of two baryons) have also been predicted. For a more detailed discussion on exotic mesons we refer to [1,2].

In recent years experimental evidence for states beyond the quark model has accumulated in the heavy quark sector. We therefore split this mini-review into three parts discussing separately light systems, heavy-light systems and heavy-heavy systems.

#### 1. Light systems

##### 1.1. Glueball candidates

Among the signatures naively expected for glueballs are (i) no free space in  $q\bar{q}$  nonets, (ii) enhanced production in gluon-rich channels such as central production and radiative  $J/\psi(1S)$  decay, (iii) decay branching fractions incompatible with SU(N) predictions for  $q\bar{q}$  states, and (iv) reduced  $\gamma\gamma$  couplings. However, mixing effects with isoscalar  $q\bar{q}$  mesons [3–10] and decay form factors [11] obscure these simple signatures.

Lattice calculations, QCD sum rules, flux tube, and constituent glue models agree that the lightest glueballs have quantum numbers  $J^{PC} = 0^{++}$  and  $2^{++}$ . Lattice calculations predict for the ground state (a  $0^{++}$  glueball) a mass around 1600 – 1700 MeV [8,12–14]) with an uncertainty of about 100 MeV, while the first excited state ( $2^{++}$ ) has a mass of about 2300 MeV. Hence, the low-mass glueballs lie in the same mass region as ordinary isoscalar  $q\bar{q}$  states, in the mass range of the  $1^3P_0(0^{++})$ ,  $2^3P_2(2^{++})$ ,  $3^3P_2(2^{++})$ , and  $1^3F_2(2^{++})$   $q\bar{q}$  states. The  $0^{-+}$  state and exotic glueballs (with non- $q\bar{q}$  quantum numbers such as  $0^{--}$ ,  $0^{+-}$ ,  $1^{-+}$ ,  $2^{-+}$ , etc.) are expected above 2 GeV [14]. The lattice calculations were performed so far in the quenched approximation. Thus neither quark loops nor mixing with conventional mesons were included, although quenching effects seem to be small [15]. For a recent comparison between quenched and unquenched lattice studies see [16].

However, one expects that glueballs will mix with nearby  $q\bar{q}$  states of the same quantum numbers. The presence of a glueball mixed with  $q\bar{q}$  would lead to a supernumerary isoscalar

state in the SU(3) classification of  $q\bar{q}$  mesons. A lattice study in full QCD (performed at unphysical quark masses corresponding to a pion mass of 400 MeV) did not identify any state with sizable overlap with pure gluonic sources [17].

In the following we focus on glueball candidates in the scalar sector. For the  $2^{++}$  sector we refer to the section on non- $q\bar{q}$  mesons in the 2006 issue of this *Review* [18], and for the  $0^{-+}$  glueball to the note on ‘The Pseudoscalar and Pseudovector Mesons in the 1400 MeV Region’ in the *Meson Listings*.

Five isoscalar resonances are well established: the very broad  $f_0(500)$  (or  $\sigma$ ), the  $f_0(980)$ , the broad  $f_0(1370)$ , and the comparatively narrow  $f_0(1500)$  and  $f_0(1710)$ , see the note on ‘Scalar Mesons below 2 GeV’ in the *Meson Listings*, and also [19]. The  $f_0(1370)$  and  $f_0(1500)$  decay mostly into pions ( $2\pi$  and  $4\pi$ ) while the  $f_0(1710)$  decays mainly into  $K\bar{K}$  final states. Naively, this suggests an  $n\bar{n}$  ( $= u\bar{u} + d\bar{d}$ ) structure for the  $f_0(1370)$  and  $f_0(1500)$ , and  $s\bar{s}$  for the  $f_0(1710)$ . The latter is not observed in  $p\bar{p}$  annihilation [20], as expected from the OZI suppression for an  $s\bar{s}$  state.

In  $\gamma\gamma$  collisions leading to  $K_S K_S$  [21] and  $K^+ K^-$  [22] a spin-0 signal is observed at the  $f_0(1710)$  mass (together with a dominant spin-2 component), while the  $f_0(1500)$  is not observed in  $\gamma\gamma \rightarrow K\bar{K}$  nor  $\pi^+\pi^-$  [23]. The  $f_0(1500)$  is also not observed by Belle in  $\gamma\gamma \rightarrow \pi^0\pi^0$ , although a shoulder is seen which could also be due to the  $f_0(1370)$  [24]. The absence of a signal in the  $\pi\pi$  channel in  $\gamma\gamma$  collisions does not favor an  $\bar{n}n$  interpretation for the  $f_0(1500)$ . The upper limit from  $\pi^+\pi^-$  excludes a large  $n\bar{n}$  content, and hence points to a mainly  $s\bar{s}$  content [25]. This is in contradiction with the small  $K\bar{K}$  decay branching ratio of the  $f_0(1500)$  [26–28]. Hence, the  $f_0(1500)$  is hard to accommodate as a  $q\bar{q}$  state. This state could be mainly glue due its absence of  $2\gamma$ -coupling, while the  $f_0(1710)$  coupling to  $2\gamma$  would be compatible with an  $s\bar{s}$  state. Indeed, Belle finds that in  $\gamma\gamma \rightarrow K_S K_S$  collisions the 1500 MeV region is dominated by the  $f'_2(1525)$ . The  $f_0(1710)$  is also observed but its production  $\times$  decay rate is too large for a glueball [29]. However, the  $2\gamma$ -couplings are sensitive to glue mixing with  $q\bar{q}$  [30].

Since the  $f_0(1370)$  does not couple strongly to  $s\bar{s}$  [28], the  $f_0(1370)$  or  $f_0(1500)$  appear to be supernumerary. The narrow width of the  $f_0(1500)$ , and its enhanced production at low transverse momentum transfer in central collisions [31–33] also favor the  $f_0(1500)$  to be non- $q\bar{q}$ . In [3] the ground state scalar nonet is made of the  $a_0(1450)$ ,  $f_0(1370)$ ,  $K_0^*(1430)$ , and  $f_0(1710)$ . The isoscalars  $f_0(1370)$  and  $f_0(1710)$  contain a small fraction of glue, while the  $f_0(1500)$  is mostly gluonic. The light scalars  $f_0(500)$ ,  $f_0(980)$ ,  $a_0(980)$ , and  $\kappa(800)$  are four-quark states or two-meson resonances (see [1] for a review). In the mixing scheme of [30], which uses central production data from WA102 and the recent hadronic  $J/\psi$  decay data from BES [34,35], glue is shared between the  $f_0(1370)$ ,  $f_0(1500)$  and  $f_0(1710)$ . The  $f_0(1370)$  is mainly  $n\bar{n}$ , the  $f_0(1500)$  mainly glue and the  $f_0(1710)$  dominantly  $s\bar{s}$ . This agrees with previous analyses [3,9], but, as already pointed out, alternative schemes

have been proposed, see, e.g. [8,36]. In particular, for a scalar glueball the two-gluon coupling to  $n\bar{n}$  appears to be suppressed by chiral symmetry [37] and therefore  $K\bar{K}$  decay could be enhanced. It was argued that chiral symmetry constraints in a multichannel analysis imply that the  $f_0(1710)$  is an unmixed scalar glueball [38], a view that is challenged [39].

The contribution of  $f_0(1500)$  production in (the supposedly gluon rich) radiative  $J/\psi$  decay is not well known. The  $f_0(1500)$  is observed by BESII in  $J/\psi \rightarrow \gamma\pi\pi$  [40] and by BESIII in  $J/\psi \rightarrow \gamma\eta\eta$  [41] with a much smaller rate than for the  $f_0(1710)$ , which speaks against a glueball interpretation for the former. However, the mass found by BES is significantly lower than the expected value. The overlap with the  $f_0(1370)$  and  $f'_2(1525)$  and the statistically limited data sample prevent a proper  $K$ -matrix analysis to be performed. Hence more data are needed in radiative  $J/\psi$  decay and in  $\gamma\gamma$  collisions to clarify the spectrum of scalar mesons.

The Dalitz plots of  $B^\pm \rightarrow \pi^\pm\pi^\pm\pi^\mp$  have been studied by BaBar [42]. A broad  $2\pi$  signal is observed around 1400 MeV which is attributed to the  $f_0(1370)$ , but could also be due to the  $f_0(1500)$ . In  $B^\pm \rightarrow K^\pm K^\pm K^\mp$  both BaBar [43] and Belle [44] observe a strong spin-0 activity in  $K\bar{K}$  around 1550 MeV.  $B^0$  decay into  $J/\psi X$  filters out the  $d\bar{d}$  content of  $X$  while  $B_s^0$  decay selects its  $s\bar{s}$  component.  $B$  decay into  $J/\psi X$  may therefore be the ideal environment to determine the flavor content of neutral mesons [45], or to distinguish  $q\bar{q}$  from tetraquark states [46]. LHCb has analyzed  $\bar{B}^0$  decay into  $J/\psi\pi^+\pi^-$  [47]. The fit to the  $\pi\pi$  mass spectrum above  $\sim 1.2$  GeV does not show any significant scalar component. However, the data analysis has been challenged [48]. For  $\bar{B}_s^0 \rightarrow J/\psi\pi^+\pi^-$  a scalar contribution from the  $f_0(1370)$  [49] or the  $f_0(1500)$  [50] is required in the 1400–1500 MeV region.

### 1.2. Tetraquark candidates and molecular bound states

The  $a_0(980)$  and  $f_0(980)$  could be tetraquark states [51–53] or  $K\bar{K}$  molecular states [54–56] due to their strong affinity for  $K\bar{K}$ , in spite of their masses being very close to threshold. For  $q\bar{q}$  states, the expected  $\gamma\gamma$  widths [57,58] are not significantly larger than for molecular states [57,59], both predictions being consistent with data. Radiative decays of the  $\phi(1020)$  into  $a_0(980)$  and  $f_0(980)$  were claimed to enable disentangling compact from molecular structures. Interpreting the data from DAPHNE [60,61] and VEPP - 2M [62,63] along the lines of [64,65] seems to favor these mesons to be tetraquark states. In [66] they are made of a four-quark core and a virtual  $K\bar{K}$  cloud at the periphery. This is challenged in [67] showing that  $\phi$  radiative decay data are consistent with a molecular structure of the light scalars. The  $f_0(980)$  is strongly produced in  $D_s^+$  decay [68]. This points to a large  $s\bar{s}$  component, assuming Cabibbo favored  $c \rightarrow s$  decay. However, the mainly  $n\bar{n}$   $f_0(1370)$  is also strongly produced in  $D_s^+$  decay, indicating that other graphs must contribute [69]. On the other hand, the contributions

from the  $f_0(500)$  and  $f_0(980)$  in the decay  $\bar{B}^0 \rightarrow J/\psi\pi^+\pi^-$  seem to exclude a tetraquark structure for these states [47], while the corresponding ones in  $\bar{B}_s^0 \rightarrow J/\psi\pi^+\pi^-$  are compatible with tetraquarks [50].

COMPASS reports a new  $1^{++}$  isovector meson at 1414 MeV, decaying into  $f_0(980)\pi$  [70]. The resonance is observed in diffractive dissociation  $\pi^-p \rightarrow \pi^-(\pi^+\pi^-)p$ . Traditionally, the  $1^{++}$  ground state nonet is believed to contain the  $a_1(1260)$ ,  $f_1(1285)$  and  $f_1(1420)$  (see the mini-review on ‘The Pseudoscalar and Pseudovector Mesons in the 1400 MeV Region’ in the *Meson Listings*). However, a molecular  $K\bar{K}\pi$  structure has been proposed for the  $f_1(1420)$  [71] in view of the proximity of the  $K^*\bar{K}$  threshold. The new  $a_1(1414)$  could then also be a molecular state, the isovector partner of the  $f_1(1420)$ .

### 1.3. Baryonia

Bound states of two nucleons have been predicted, but have remained elusive. The  $f_2(1565)$  which is only observed in  $\bar{p}p$  annihilation [72,73] is a good candidate for a  $2^{++}$   $\bar{p}p$  bound state. Enhancements in the  $\bar{p}p$  mass spectrum have also been reported below  $\bar{p}p$  threshold, in  $J/\psi \rightarrow \gamma\bar{p}p$  [74–76] and in  $B^+ \rightarrow K^+\bar{p}p$ ,  $B^0 \rightarrow K_S^0\bar{p}p$  [77,78] and  $\bar{B}^0 \rightarrow D^0\bar{p}p$  [79]. This enhancement could be due to a  $0^{-+}$  baryonium [80] but other explanations have been proposed, such as the dynamics of the fragmentation mechanism [78] or the attractive  $^1S_0$  ( $\bar{p}p$ ) wave [81–83].

The pronounced signal observed in  $e^+e^- \rightarrow \Lambda_c^+\Lambda_c^-$  around  $\sqrt{s} = 4.63$  GeV by Belle [84] was argued to be a strong evidence in favor of an interpretation of  $Y(4660)$  as charmed baryonium [85]. However, this picture was challenged in [86].

### 1.4. Hybrid mesons

Hybrids may be viewed as  $q\bar{q}$  mesons with a vibrating gluon flux tube. In contrast to glueballs, they can have isospin 0 or 1. The mass spectrum of hybrids with exotic (non- $q\bar{q}$ ) quantum numbers was predicted in [87], while [88] also deals with non-exotic quantum numbers. The ground state hybrids with quantum numbers ( $0^{-+}$ ,  $1^{-+}$ ,  $1^{--}$ , and  $2^{-+}$ ) are expected around 1.7 to 1.9 GeV. Lattice calculations predict that the hybrid with exotic quantum numbers  $1^{-+}$  lies at a mass of  $1.9 \pm 0.2$  GeV [89,90]. Most hybrids are expected to be rather broad, but some can be as narrow as 100 MeV [91]. They prefer to decay into a pair of  $S$ - and  $P$ -wave mesons. The lattice study in [17], based on full QCD with pion masses around 400 MeV, finds that several of the high-lying states observed in their spectrum show significant overlap with gluon rich source terms.

A  $J^{PC} = 1^{-+}$  exotic meson,  $\pi_1(1400)$ , was reported in  $\pi^-p \rightarrow \eta\pi^-p$  [92,93] and in  $\pi^-p \rightarrow \eta\pi^0n$  [94]. It was observed as an interference between the angular momentum  $L = 1$  and  $L = 2$   $\eta\pi$  amplitudes, leading to a forward/backward asymmetry in the  $\eta\pi$  angular distribution. This state has been reported earlier in  $\pi^-p$  reactions [95], but ambiguous solutions

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in the partial wave analysis were pointed out in [96,97]. A resonating  $1^{-+}$  contribution to the  $\eta\pi$   $P$ -wave is also required in the Dalitz plot analysis of  $\bar{p}n$  annihilation into  $\pi^-\pi^0\eta$  [98], and in  $\bar{p}p$  annihilation into  $\pi^0\pi^0\eta$  [99]. Mass and width are consistent with the results of Ref. [92].

Another  $1^{-+}$  state,  $\pi_1(1600)$ , decaying into  $\rho\pi$ , was reported by COMPASS with 190 GeV pions hitting a lead target [100]. It was observed earlier in  $\pi^-p$  interactions in the decay modes  $\eta'\pi$  [101],  $f_1(1285)\pi$  [102], and  $\omega\pi\pi$  [103],  $b_1(1235)\pi$ , but not  $\eta\pi$  [104]. A strong enhancement in the  $1^{-+}$   $\eta'\pi$  wave, compared to  $\eta\pi$ , was reported at this mass in [105]. Ref. [106] suggests that a Deck-generated  $\eta\pi$  background from final state rescattering in  $\pi_1(1600)$  decay could mimic  $\pi_1(1400)$ . However, this mechanism is absent in  $\bar{p}p$  annihilation. The  $\eta\pi\pi$  data require  $\pi_1(1400)$  and cannot accommodate a state at 1600 MeV [107]. Finally, evidence for a  $\pi_1(2015)$  has also been reported [102,103].

The flux tube model and the lattice concur to predict a hybrid mass of about 1.9 GeV while the  $\pi_1(1400)$  and  $\pi_1(1600)$  are lighter. As isovectors,  $\pi_1(1400)$  and  $\pi_1(1600)$  cannot be glueballs. The coupling to  $\eta\pi$  of the former points to a four-quark state [108], while the strong  $\eta'\pi$  coupling of the latter is favored for hybrid states [109,110]. The mass of  $\pi_1(1600)$  is also not far below the lattice prediction.

Hybrid candidates with  $J^{PC} = 0^{-+}$ ,  $1^{--}$ , and  $2^{-+}$  have also been reported. The  $\pi(1800)$  decays mostly to a pair of  $S$ - and  $P$ -wave mesons [100,111], in line with expectations for  $0^{-+}$  hybrid mesons. This meson is also somewhat narrow if interpreted as the second radial excitation of the pion. The evidence for  $1^{--}$  hybrids required in  $e^+e^-$  annihilation and in  $\tau$  decays has been discussed in Ref. [112]. A candidate for the  $2^{-+}$  hybrid, the  $\eta_2(1870)$ , was reported in  $\gamma\gamma$  interactions [113], in  $\bar{p}p$  annihilation [114], and in central production [115]. The near degeneracy of  $\eta_2(1645)$  and  $\pi_2(1670)$  suggests ideal mixing in the  $2^{-+}$   $q\bar{q}$  nonet, and hence, the second isoscalar should be mainly  $s\bar{s}$ . However,  $\eta_2(1870)$  decays mainly to  $a_2(1320)\pi$  and  $f_2(1270)\pi$  [114], with a relative rate compatible with a hybrid state [88].

## 2. Heavy-light systems

Two very narrow states,  $D_{s0}^*(2317)^\pm$  and  $D_{s1}(2460)^\pm$ , were observed at B factories [116,117]. They lie far below the predicted masses for the two expected broad  $P$ -wave  $c\bar{s}$  mesons. These states have hence been interpreted as four-quark states [118–120] or  $DK$  ( $DK^*$ ) molecules [121–123]. However, strong cusp effects, due to the nearby  $DK$  ( $DK^*$ ) thresholds, could shift their masses downwards and quench the observed widths, an effect similar to that occurring for the  $a_0(980)$  and  $f_0(980)$  mesons, which lie just below  $K\bar{K}$  threshold. A hadronic width of typically 100 keV would be the unequivocal signature for the molecular nature of  $D_{s0}^*(2317)^\pm$  [123,124]. Its measured width is currently less than 3.8 MeV.

## 3. Heavy-heavy systems

A search for multiquark states containing a  $c$  (or a  $b$ ) quark is promising since the charmonium spectrum can be predicted accurately, and because those mesons lying below the  $D\bar{D}$  or  $D\bar{D}^*$  thresholds should be narrow. Several states have been observed recently in the charmonium region. With the discovery of the  $X(3872)$  in  $B^\pm \rightarrow K^\pm X$  ( $X \rightarrow J/\psi \pi^+\pi^-$ ) by Belle [126], soon confirmed by BaBar [127], many searches for states beyond the standard quark model were initiated both in the charm and in the bottom sectors. In the decay  $\Lambda_b^0 \rightarrow J/\psi K^- p$  the LHCb collaboration has recently reported the observation of two new baryons decaying into  $J/\psi p$ , which are candidates for heavy pentaquark states [128]. For an updated collection of the currently available experimental information on multiquark states we refer to the mini-review on ‘Spectroscopy of mesons containing two heavy quarks’ in the *Meson Listings*.

When restricting ourselves to confirmed states we are faced with eight states that do not seem to fit into the standard quark model. This is clear for the four established charged states ( $Z_c(3900)^\pm$  and  $Z_c(4430)^\pm$  in the charmonium sector, and  $Z_b(10610)^\pm$  and  $Z_b(10650)^\pm$  in the bottomonium sector). The neutral ones ( $X(3872)$ ,  $Y(4260)$ ,  $Y(4360)$ ,  $Y(4660)$ ) also challenge the standard quark model since their masses and decay properties are in conflict with expectations.

The quantum numbers of the  $X(3872)$  have been determined by LHCb to be  $J^{PC} = 1^{++}$ , first by assuming the angular momentum zero between the  $J/\psi$  and the dipion [129] and then by relaxing this constraint [130]. The  $X(3872)$  can hardly be identified with the  $2^3P_1$   $\chi'_{c1}$  since the latter is predicted to lie about 100 MeV higher in mass [131]. Instead, the  $X(3940)$  reported by Belle in  $e^+e^- \rightarrow J/\psi X$ , decaying into  $D^*\bar{D}$  but not into  $D\bar{D}$  [132], and also observed in  $B \rightarrow K(X \rightarrow \omega J/\psi)$  [133] could be the  $\chi'_{c1}$ . The  $2^3P_2$  tensor partner ( $\chi'_{c2}$ ) was reported by Belle at 3931 MeV in  $\gamma\gamma$  interactions [134].

The  $X(3872)$  lies close to the  $D^0\bar{D}^{*0}$  threshold and therefore the most natural explanation for this state is a  $1^{++}$   $D\bar{D}^*$  molecule [135] for which strong isospin breaking is predicted [135,136] due to the nearby  $D^+D^{*-}$  threshold. Indeed, the comparable rates for  $\omega J/\psi$  and  $\rho^0 J/\psi$  are consistent with an interpretation of  $X(3872)$  as an isoscalar  $D\bar{D}^*$  molecule when the different widths of the  $\rho$  and  $\omega$  are taken into account [137]. A four-quark state  $cq\bar{c}\bar{q}'$  is also possible [120] but unlikely, since the charged partner of the  $X(3872)$  has not been observed (e.g. in  $B^- \rightarrow \bar{K}^0 X^-$  nor in  $B^0 \rightarrow K^+ X^-$ , where  $X^- \rightarrow J/\psi \pi^-\pi^0$  [138]). The claim that  $X(3872)$  must be a compact (tetraquark) state, since it is also produced at very high  $p_T$  in  $\bar{p}p$  collisions [139], was challenged in [140] which stresses the importance of rescattering, see also Ref. [141].

A broad structure,  $Y(4260)$ , decaying into  $J/\psi \pi^+\pi^-$  was reported by BaBar in initial state radiation  $e^+e^- \rightarrow \gamma(e^+e^- \rightarrow Y(4260))$  [142]. A charmonium state with the quantum numbers  $1^{--}$  is not expected in this mass region. In addition,



See key on page 601

a charmonium at this mass should have a significant coupling to  $\bar{D}D$ , a decay channel that is not observed for the  $Y(4260)$ . This state could be a hybrid charmonium with a spin-1  $\bar{c}c$  core [143,144]. However, provided that the observation of  $Y(4260)$  decay into  $h_c(1P)\pi\pi$  by BESIII [145] is confirmed, the hybrid hypothesis would be under pressure, since the spin of the heavy quarks (coupled to zero in the  $h_c(1P)$ ) should be conserved in leading order in the expansion in  $(\Lambda_{\text{QCD}}/m_c)$ . (The individual conservation of the heavy quark spin and the total angular momentum of the light quark cloud is a consequence of the heavy quark spin symmetry, see the review on ‘Heavy-Quark and Soft-Collinear Effective Theory’ in this issue of the *Review*.)

The same criticism applies to the hadrocharmonium interpretation of the  $Y(4260)$  which describes this state as spin-1 quarkonium surrounded by a light quark cloud [146]. To circumvent the spin symmetry argument Ref. [147] argues that  $Y(4260)$  and  $Y(4360)$  could be mixtures of two hadrocharmonia with spin-triplet and spin-singlet heavy quark pairs. The same kind of mixing could also operate for a hybrid.

A dominant  $D_1\bar{D}$  component in the  $Y(4260)$  [148] would explain naturally why  $Z_c(3900)^\pm$  (interpreted by the authors as a  $\bar{D}D^*$  bound state) is seen in  $Y(4260) \rightarrow \pi^\mp Z_c(3900)^\pm$ . A prominent  $D_1\bar{D}$  component in the  $Y(4260)$  would in addition explain the copious production of  $X(3872)$  in  $Y(4260)$  radiative decays [149]. The  $Y(4360)$  as  $D_1\bar{D}^*$  bound state could be the spin partner of the  $Y(4260)$  [150,151], but a detailed microscopic calculation is still lacking.

The tetraquark picture explains the observed  $Y$  states and, when including a tailor-made spin-spin interaction, is also capable to describe both  $Z_c(3900)^{\pm,0}$  and  $Z_c(4020)^\pm$  [152], and even the recently confirmed  $Z(4430)^\pm$  by Belle [153]. However, the model predicts many additional charged and neutral states which have not yet been discovered.

The charged states  $Z_c(3900)^\pm$ , first observed by BESIII [154] and the to be confirmed  $Z_c(4020)^\pm$  [155] decay predominantly into  $\bar{D}D^*$  and  $\bar{D}^*D^*$ , respectively, while  $Z_b(10610)^{\pm,0}$  and  $Z_b(10650)^\pm$  [156,157] decay predominantly into  $\bar{B}B^*$  and  $\bar{B}^*B^*$ , respectively, although all of them were discovered in the decay mode heavy quarkonium and pion. This suggests that the states are close relatives and their interactions are connected via heavy quark flavor symmetry. A molecular interpretation for the bottomonium states was proposed shortly after the discovery of the  $Z_b^\pm$  states [158] and also shortly after that of the  $Z_c(3900)^\pm$  [148]. However, their properties also appear to be consistent with tetraquark structures [159].

The heaviest confirmed charged state in the charmonium sector is the  $Z(4430)^\pm$  observed by Belle [153]. It is interpreted as hadrocharmonium [146],  $\bar{D}_1D^*$  molecule [160] as well as tetraquark state [152]. Alternatively, in [161] the  $Z(4430)^\pm$  can be explained as a cross-channel effect.

It should be stressed that the various scenarios, while describing the data, also make decisive predictions, e.g. for other decay channels. The forthcoming data on heavy meson spectroscopy from various facilities should soon provide a much

deeper understanding on how QCD forms matter out of quarks and gluons.

## References

1. C. Amsler and N.A. Törnqvist, Phys. Reports **389**, 61 (2004).
2. N. Brambilla *et al.*, Eur. Phys. J. **C74**, 2981 (2014).
3. C. Amsler and F.E. Close, Phys. Rev. **D53**, 295 (1996).
4. N.A. Törnqvist, Phys. Rev. Lett. **76**, 1575 (1996).
5. A.V. Anisovich and A.V. Sarantsev, Phys. Lett. **B395**, 123 (1997).
6. M. Boglione *et al.*, Phys. Rev. Lett. **79**, 1998 (1997).
7. P. Minkowski and W. Ochs, Eur. Phys. J. **C9**, 283 (1999).
8. W. Lee and D. Weingarten, Phys. Rev. **D61**, 014015 (2000).
9. F.E. Close and A. Kirk, Eur. Phys. J. **C21**, 531 (2001).
10. H.Y. Cheng *et al.*, Phys. Rev. **D92**, 094006 (2015), see also Phys. Rev. **D74**, 094005 (2006).
11. T. Barnes *et al.*, Phys. Rev. **D55**, 4157 (1997).
12. G.S. Bali *et al.*, Phys. Lett. **B309**, 378 (1993).
13. C.J. Morningstar and M.J. Peardon, Phys. Rev. **D56**, 4043 (1997).
14. Y. Chen *et al.*, Phys. Rev. **D73**, 014516 (2006).
15. C.M. Richards *et al.*, Phys. Rev. **D82**, 034501 (2010).
16. E. Gregory *et al.*, JHEP **1210**, 170 (2012).
17. J.J. Dudek *et al.*, Phys. Rev. **D83**, 111502 (2011).
18. W.-M. Yao *et al.*, J. Phys. **G33**, 1 (2006) p.949.
19. C. Amsler, Rev. Mod. Phys. **70**, 1293 (1998).
20. C. Amsler *et al.*, Eur. Phys. J. **C23**, 29 (2002).
21. M. Acciarri *et al.*, Phys. Lett. **B501**, 173 (2001).
22. K. Abe *et al.*, Eur. Phys. J. **C32**, 323 (2004).
23. R. Barate *et al.*, Phys. Lett. **B472**, 189 (2000).
24. S. Uehara *et al.*, Phys. Rev. **D78**, 052004 (2008).
25. C. Amsler, Phys. Lett. **B541**, 22 (2002).
26. A. Abele *et al.*, Phys. Lett. **B385**, 425 ((1996).
27. A. Abele *et al.*, Phys. Rev. **D57**, 3860 (1998).
28. D. Barberis *et al.*, Phys. Lett. **B462**, 462 (1999).
29. S. Uehara *et al.*, Prog. Theor. Exp. Phys. **2013**, 123C01 (2013).
30. F.E. Close and Q. Zhao, Phys. Rev. **D71**, 094022 (2005).
31. F.E. Close *et al.*, Phys. Lett. **B397**, 333 (1997).
32. F.E. Close, Phys. Lett. **B419**, 387 (1998).
33. A. Kirk, Phys. Lett. **B489**, 29 (2000).
34. M. Ablikim *et al.*, Phys. Lett. **B603**, 138 (2004).
35. M. Ablikim *et al.*, Phys. Lett. **B607**, 243 (2005).
36. J. Chen *et al.*, Mod. Phys. Lett. **A24**, 1517 (2009).
37. M. Chanowitz, Phys. Rev. Lett. **95**, 172001 (2005).
38. M. Albaladejo and J.A. Oller, Phys. Rev. Lett. **101**, 252002 (2008).
39. L.S. Geng and E. Oset, Phys. Rev. **D79**, 074009 (2009).
40. M. Ablikim *et al.*, Phys. Lett. **B642**, 441 (2006).
41. M. Ablikim *et al.*, Phys. Rev. **D87**, 092009 (2013).
42. B. Aubert *et al.*, Phys. Rev. **D79**, 072006 (2009).
43. B. Aubert *et al.*, Phys. Rev. **D74**, 032003 (2006).
44. A. Garmash *et al.*, Phys. Rev. **D71**, 092003 (2005).
45. C.D. Lü *et al.*, Eur. Phys. J. **A49**, 58 (2013).

## Meson Particle Listings

Non- $q\bar{q}$  Candidates

46. S. Stone and L. Zhang, Phys. Rev. Lett. **111**, 062001 (2013).
47. R. Aaij *et al.*, Phys. Rev. **D90**, 012003 (2014).
48. F.E. Close and A. Kirk, Phys. Rev. **D91**, 114015 (2015).
49. R. Aaij *et al.*, Phys. Rev. **D86**, 052006 (2012).
50. R. Aaij *et al.*, Phys. Rev. **D89**, 092006 (2014).
51. R.L. Jaffe, Phys. Rev. **D15**, 267 (1977), Phys. Rev. **D15**, 281 (1977).
52. M. Alford and R.L. Jaffe, Nucl. Phys. **B578**, 367 (2000).
53. G. 't Hooft *et al.*, Phys. Lett. **B662**, 424 (2008).
54. J. Weinstein and N. Isgur, Phys. Rev. **D41**, 2236 (1990).
55. G. Janssen *et al.*, Phys. Rev. **D52**, 2690 (1995).
56. M.P. Locher *et al.*, Eur. Phys. J. **C4**, 317 (1998).
57. J.A. Oller and E. Oset, Hadron 97 Conf. AIP Conf. Proc. **432**, 413 (1997).
58. R. Delbourgo *et al.*, Phys. Lett. **B446**, 332 (1999).
59. C. Hanhart *et al.*, Phys. Rev. **D75**, 074015 (2007).
60. A. Aloisio *et al.*, Phys. Lett. **B536**, 209 (2002).
61. A. Aloisio *et al.*, Phys. Lett. **B537**, 21 (2002).
62. R.R. Akhmetshin *et al.*, Phys. Lett. **B462**, 371 (1999).
63. M.N. Achasov *et al.*, Phys. Lett. **B479**, 53 (2000).
64. F.E. Close *et al.*, Nucl. Phys. **B389**, 513 (1993).
65. N.N. Achasov *et al.*, Phys. Rev. **D56**, 203 (1997).
66. F.E. Close and N. Törnqvist, J. Phys. **G28**, R249 (2002).
67. Y.S. Kalashnikova *et al.*, Eur. Phys. J. **A24**, 437 (2005).
68. E.M. Aitala *et al.*, Phys. Rev. Lett. **86**, 765 (2001).
69. H.Y. Cheng, Phys. Rev. **D67**, 054021 (1997).
70. C. Adolph *et al.*, Phys. Rev. Lett. **115**, 082001 (2015).
71. R.S. Longacre, Phys. Rev. **D42**, 874 (1990).
72. B. May *et al.*, Z. Phys. **C46**, 203 (1990).
73. A. Bertin *et al.*, Phys. Rev. **D57**, 55 (1998).
74. J.Z. Bai *et al.*, Phys. Rev. Lett. **91**, 022001 (2003).
75. J.P. Alexander *et al.*, Phys. Rev. **D82**, 092002 (2010).
76. M. Ablikim *et al.*, Phys. Rev. Lett. **108**, 112003 (2012).
77. K. Abe *et al.*, Phys. Rev. Lett. **88**, 181803 (2002).
78. M.-Z. Wang *et al.*, Phys. Lett. **B617**, 141 (2005).
79. K. Abe *et al.*, Phys. Rev. Lett. **89**, 151802 (2002).
80. G.-J. Ding and M.L. Yan, Phys. Rev. **C72**, 015208 (2005).
81. B. Loiseau and S. Wycech, Phys. Rev. **C72**, 011001 (2005).
82. A. Sibirtsev *et al.*, Phys. Rev. **D71**, 054010 (2005).
83. X. W. Kang, J. Haidenbauer, and U.-G. Meissner, Phys. Rev. **D91**, 074003 (2015); see also J. Haidenbauer, X.-W. Kang, and U.-G. Meissner, Nucl. Phys. **A929**, 102 (2014).
84. G. Pakhlova *et al.*, Phys. Rev. Lett. **101**, 172001 (2008).
85. G. Cotugno *et al.*, Phys. Rev. Lett. **104**, 132005 (2010).
86. F.-K. Guo *et al.*, Phys. Rev. **D82**, 094008 (2010).
87. N. Isgur *et al.*, Phys. Rev. Lett. **54**, 869 (1985).
88. F.E. Close and P.R. Page, Nucl. Phys. **B433**, 233 (1995).
89. P. Lacock *et al.*, Phys. Lett. **B401**, 308 (1997).
90. C. Bernard *et al.*, Phys. Rev. **D56**, 7039 (1997).
91. P.R. Page *et al.*, Phys. Rev. **D59**, 034016 (1999).
92. D.R. Thompson *et al.*, Phys. Rev. Lett. **79**, 1630 (1997).
93. S.U. Chung *et al.*, Phys. Rev. **D60**, 092001 (1999).
94. G.S. Adams *et al.*, Phys. Lett. **B657**, 27 (2007).
95. D.M. Alde *et al.*, Phys. Lett. **B205**, 397 (1988).
96. Y.D. Prokoshkin and S.A. Sadovsky, Phys. Atom. Nucl. **58**, 606 (1995).
97. Y.D. Prokoshkin and S.A. Sadovsky, Phys. Atom. Nucl. **58**, 853 (1995).
98. A. Abele *et al.*, Phys. Lett. **B423**, 175 (1998).
99. A. Abele *et al.*, Phys. Lett. **B446**, 349 (1999).
100. M.G. Alekseev *et al.*, Phys. Rev. Lett. **104**, 241803 (2010).
101. E.I. Ivanov *et al.*, Phys. Rev. Lett. **86**, 3977 (2001).
102. J. Kuhn *et al.*, Phys. Lett. **B595**, 109 (2004).
103. M. Lu *et al.*, Phys. Rev. Lett. **94**, 032002 (2005).
104. Yu.P. Gouz *et al.*, Proc. XXVI Int. Conf. on HEP, Dallas (1992).
105. G.M. Beladidze *et al.*, Phys. Lett. **B313**, 276 (1993).
106. A. Donnachie *et al.*, Phys. Rev. **D58**, 114012 (1998).
107. W. Dünnweber, Nucl. Phys. **A663-664**, 592C (1999).
108. S.U. Chung *et al.*, Eur. Phys. J. **A15**, 539 (2002).
109. F.E. Close and H.J. Lipkin, Phys. Lett. **B196**, 245 (1987).
110. F. Iddir and A.S. Safir, Phys. Lett. **B507**, 183 (2001).
111. D.V. Amelin *et al.*, Phys. Lett. **B356**, 595 (1995).
112. A. Donnachie and Yu.S. Kalashnikova, Phys. Rev. **D60**, 114011 (1999).
113. K. Karch *et al.*, Z. Phys. **C54**, 33 (1992).
114. J. Adomeit *et al.*, Z. Phys. **C71**, 227 (1996).
115. D. Barberis *et al.*, Phys. Lett. **B413**, 217 (1997).
116. B. Aubert *et al.*, Phys. Rev. Lett. **90**, 242001 (2003).
117. D. Besson *et al.*, Phys. Rev. **D68**, 032002 (2003).
118. H.-Y. Cheng and W.-S. Hou, Phys. Lett. **B566**, 193 (2003).
119. K. Terasaki, Phys. Rev. **D68**, 011501 (2003).
120. L. Maiani *et al.*, Phys. Rev. **D71**, 014028 (2005).
121. T. Barnes *et al.*, Phys. Rev. **D68**, 054006 (2003).
122. E.E. Kolomeitsev and M.F.M. Lutz, Phys. Lett. **B582**, 39 (2004).
123. L. Liu *et al.*, Phys. Rev. **D87**, 014508 (2013).
124. A. Faessler *et al.*, Phys. Rev. **D76**, 014005 (2007).
125. M.F.M. Lutz and M. Soyeur, Nucl. Phys. **A813**, 14 (2008).
126. S.-K. Choi *et al.*, Phys. Rev. Lett. **91**, 262001 (2003).
127. B. Aubert *et al.*, Phys. Rev. **D71**, 071103R (2005).
128. R. Aaij *et al.*, Phys. Rev. Lett. **115**, 072001 (2015).
129. R. Aaij *et al.*, Phys. Rev. Lett. **110**, 222001 (2013).
130. R. Aaij *et al.*, Phys. Rev. **D92**, 011102 (2015).
131. T. Barnes and S. Godfrey, Phys. Rev. **D69**, 054008 (2004).
132. K. Abe *et al.*, Phys. Rev. Lett. **98**, 082001 (2007).
133. S.-K. Choi *et al.*, Phys. Rev. Lett. **94**, 182002 (2005).
134. S. Uehara *et al.*, Phys. Rev. Lett. **96**, 082003 (2006).
135. N. Törnqvist, Phys. Lett. **B590**, 209 (2004).
136. E. Swanson, Phys. Lett. **B588**, 189 (2004).
137. D. Gamermann and E. Oset, Phys. Rev. **D80**, 014003 (2009).
138. B. Aubert *et al.*, Phys. Rev. **D71**, 031501R (2005).
139. C. Bignamini *et al.*, Phys. Rev. Lett. **103**, 162001 (2009).

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# Meson Particle Listings

## Non- $q\bar{q}$ Candidates

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- |   |  |
|---|--|
| <p>140. P. Artoisenet and E. Braaten, Phys. Rev. <b>D81</b>, 114018 (2010).</p> <p>141. F.-K. Guo <i>et al.</i>, JHEP <b>1405</b>, 138 (2014).</p> <p>142. B. Aubert <i>et al.</i>, Phys. Rev. Lett. <b>95</b>, 142001 (2005).</p> <p>143. F.E. Close and P.R. Page, Phys. Lett. <b>B628</b>, 215 (2005).</p> <p>144. E. Kou, Phys. Lett. <b>B631</b>, 164 (2005).</p> <p>145. M. Ablikim <i>et al.</i>, Phys. Rev. Lett. <b>111</b>, 242001 (2013).</p> <p>146. M.B. Voloshin, Prog. in Part. Nucl. Phys. <b>61</b>, 455 (2008).</p> <p>147. X. Li, M.B. Voloshin, Mod. Phys. Lett. <b>A29</b>, 50060 (2014).</p> <p>148. Q. Wang <i>et al.</i>, Phys. Rev. Lett. <b>111</b>, 132003 (2013).</p> <p>149. F.K. Guo <i>et al.</i>, Phys. Lett. <b>B725</b>, 127 (2013).</p> <p>150. Q. Wang <i>et al.</i>, Phys. Rev. <b>D89</b>, 034001 (2014).</p> | <p>151. L. Ma <i>et al.</i>, Phys. Rev. <b>D91</b>, 034002 (2015).</p> <p>152. L. Maiani <i>et al.</i>, Phys. Rev. <b>D89</b>, 114010 (2014).</p> <p>153. B. Chilikin <i>et al.</i>, Phys. Rev. <b>D88</b>, 074026 (2013).</p> <p>154. M. Ablikim <i>et al.</i>, Phys. Rev. Lett. <b>110</b>, 252001 (2013).</p> <p>155. M. Ablikim <i>et al.</i>, Phys. Rev. Lett. <b>111</b>, 242002 (2013).</p> <p>156. P. Krokovny <i>et al.</i>, Phys. Rev. <b>D88</b>, 052016 (2013).</p> <p>157. A. Bondar <i>et al.</i>, Phys. Rev. Lett. <b>108</b>, 122001 (2012).</p> <p>158. A. Bondar <i>et al.</i>, Phys. Rev. <b>D84</b>, 054010 (2011).</p> <p>159. A. Ali <i>et al.</i>, Phys. Rev. <b>D91</b>, 017502 (2015).</p> <p>160. T. Branz <i>et al.</i>, Phys. Rev. <b>D82</b>, 054025 (2010).</p> <p>161. P. Pakhlov, Phys. Lett. <b>B702</b>, 139 (2011);<br/>P. Pakhlov and T. Uglov, Phys. Lett. <b>B748</b>, 183 (2015).</p> |
|---|--|
-



**$N$  BARYONS ( $S = 0, I = 1/2$ )**

$p$	1503
$n$	1512
$N$ resonances	1523

 **$\Delta$  BARYONS ( $S = 0, I = 3/2$ )**

$\Delta$ resonances	1554
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 **$\Lambda$  BARYONS ( $S = -1, I = 0$ )**

$\Lambda$	1574
$\Lambda$ resonances	1578

 **$\Sigma$  BARYONS ( $S = -1, I = 1$ )**

$\Sigma^+$	1594
$\Sigma^0$	1596
$\Sigma^-$	1596
$\Sigma$ resonances	1599

 **$\Xi$  BARYONS ( $S = -2, I = 1/2$ )**

$\Xi^0$	1620
$\Xi^-$	1622
$\Xi$ resonances	1625

 **$\Omega$  BARYONS ( $S = -3, I = 0$ )**

$\Omega^-$	1632
$\Omega$ resonances	1633

**CHARMED BARYONS ( $C = +1$ )**

$\Lambda_c^+$	1637
$\Lambda_c(2595)^+$	1643
$\Lambda_c(2625)^+$	1643
$\Lambda_c(2765)^+$	1644
$\Lambda_c(2880)^+$	1644
$\Lambda_c(2940)^+$	1645
$\Sigma_c(2455)$	1645
$\Sigma_c(2520)$	1646
$\Sigma_c(2800)$	1647
$\Xi_c^+$	1648
$\Xi_c^0$	1649
$\Xi_c^{\prime+}$	1650
$\Xi_c^{\prime0}$	1651
$\Xi_c(2645)$	1651
$\Xi_c(2790)$	1651
$\Xi_c(2815)$	1651
$\Xi_c(2930)$	1652
$\Xi_c(2970)$	1652
$\Xi_c(3055)$	1653
$\Xi_c(3080)$	1653
$\Xi_c(3123)$	1653
$\Omega_c^0$	1653
$\Omega_c(2770)^0$	1654

**DOUBLY-CHARMED BARYONS ( $C = +2$ )**

$\Xi_{cc}^+$	1655
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**BOTTOM (BEAUTY) BARYONS ( $B = -1$ )**

$\Lambda_b^0$	1656
$\Lambda_b(5912)^0$	1660
$\Lambda_b(5920)^0$	1661
$\Sigma_b$	1661
$\Sigma_b^*$	1661
$\Xi_b^0, \Xi_b^-$	1662
$\Xi_b^{\prime0}(5935)^-$	1663
$\Xi_b(5945)^0$	1663
$\Xi_b^*(5955)^-$	1664
$\Omega_b^-$	1664
$b$ -baryon ADMIXTURE ( $\Lambda_b, \Xi_b, \Sigma_b, \Omega_b$ )	1665

**EXOTIC BARYONS**

$P_c(4380)^+$	1671
$P_c(4450)^+$	1671

**Notes in the Baryon Listings**

Baryon decay parameters	1515
$N$ and $\Delta$ resonances (rev.)	1518
Baryon magnetic moments	1574
$\Lambda$ and $\Sigma$ resonances	1577
Pole structure of the $\Lambda(1405)$ region (new)	1578
Radiative hyperon decays	1621
$\Xi$ resonances	1625
Charmed baryons	1635
Pentaquarks (new)	1667



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# Baryon Particle Listings

p

## N BARYONS (S = 0, I = 1/2)

$$p, N^+ = uud; \quad n, N^0 = udd$$



$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+) \text{ Status: } ***$$

### p MASS (atomic mass units u)

The mass is known much more precisely in u (atomic mass units) than in MeV. See the next data block.

VALUE (u)	DOCUMENT ID	TECN	COMMENT
<b>1.007276466879 ± 0.000000000091</b>	MOHR 16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.007276466812 ± 0.000000000090	MOHR 12	RVUE	2010 CODATA value
1.00727646677 ± 0.000000000010	MOHR 08	RVUE	2006 CODATA value
1.00727646688 ± 0.000000000013	MOHR 05	RVUE	2002 CODATA value
1.00727646688 ± 0.000000000013	MOHR 99	RVUE	1998 CODATA value
1.007276470 ± 0.0000000012	COHEN 87	RVUE	1986 CODATA value

### p MASS (MeV)

The mass is known much more precisely in u (atomic mass units) than in MeV. The conversion from u to MeV,  $1 \text{ u} = 931.494 \text{ 0054(57) MeV}/c^2$  (MOHR 16, the 2014 CODATA value), involves the relatively poorly known electronic charge.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>938.2720813 ± 0.0000058</b>	MOHR 16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
938.272046 ± 0.000021	MOHR 12	RVUE	2010 CODATA value
938.272013 ± 0.000023	MOHR 08	RVUE	2006 CODATA value
938.272029 ± 0.000080	MOHR 05	RVUE	2002 CODATA value
938.271998 ± 0.000038	MOHR 99	RVUE	1998 CODATA value
938.27231 ± 0.00028	COHEN 87	RVUE	1986 CODATA value
938.2796 ± 0.0027	COHEN 73	RVUE	1973 CODATA value

$$|m_p - m_{\bar{p}}|/m_p$$

A test of *CPT* invariance. Note that the comparison of the  $\bar{p}$  and  $p$  charge-to-mass ratio, given in the next data block, is much better determined.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7 × 10<sup>-10</sup></b>	90	<sup>1</sup> HORI	11	SPEC $\bar{p}e^-$ -He atom
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2 × 10 <sup>-9</sup>	90	<sup>1</sup> HORI	06	SPEC $\bar{p}e^-$ -He atom
<1.0 × 10 <sup>-8</sup>	90	<sup>1</sup> HORI	03	SPEC $\bar{p}e^-$ <sup>4</sup> He, $\bar{p}e^-$ <sup>3</sup> He
<6 × 10 <sup>-8</sup>	90	<sup>1</sup> HORI	01	SPEC $\bar{p}e^-$ -He atom
<5 × 10 <sup>-7</sup>		<sup>2</sup> TORII	99	SPEC $\bar{p}e^-$ -He atom

<sup>1</sup> HORI 01, HORI 03, HORI 06, and HORI 11 use the more-precisely-known constraint on the  $\bar{p}$  charge-to-mass ratio of GABRIELSE 99 (see below) to get their results. Their results are not independent of the HORI 01, HORI 03, HORI 06, and HORI 11 values for  $|q_p + q_{\bar{p}}|/e$ , below.

<sup>2</sup> TORII 99 uses the more-precisely-known constraint on the  $\bar{p}$  charge-to-mass ratio of GABRIELSE 95 (see below) to get this result. This is not independent of the TORII 99 value for  $|q_p + q_{\bar{p}}|/e$ , below.

### $\bar{p}/p$ CHARGE-TO-MASS RATIO, $|q_{\bar{p}}/m_{\bar{p}}|/(q_p/m_p)$

A test of *CPT* invariance. Listed here are measurements involving the *inertial* masses. For a discussion of what may be inferred about the ratio of  $\bar{p}$  and  $p$  *gravitational* masses, see ERICSON 90; they obtain an upper bound of  $10^{-6}$ – $10^{-7}$  for violation of the equivalence principle for  $\bar{p}$ 's.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.99999999991 ± 0.00000000009</b>	GABRIELSE 99	TRAP	Penning trap
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.0000000015 ± 0.0000000011	<sup>1</sup> GABRIELSE 95	TRAP	Penning trap
1.0000000023 ± 0.0000000042	<sup>2</sup> GABRIELSE 90	TRAP	Penning trap

<sup>1</sup> Equation (2) of GABRIELSE 95 should read  $M(\bar{p})/M(p) = 0.999\,999\,9985(11)$  (G. Gabrielse, private communication).

<sup>2</sup> GABRIELSE 90 also measures  $m_{\bar{p}}/m_{e^-} = 1836.152660 \pm 0.000083$  and  $m_p/m_{e^-} = 1836.152680 \pm 0.000088$ . Both are completely consistent with the 1986 CODATA (COHEN 87) value for  $m_p/m_{e^-}$  of  $1836.152701 \pm 0.000037$ .

$$(|q_{\bar{p}}/m_{\bar{p}} - q_p/m_p)/q_p/m_p$$

A test of *CPT* invariance. Taken from the  $\bar{p}/p$  charge-to-mass ratio, above.

VALUE	DOCUMENT ID
<b>(−9 ± 9) × 10<sup>-11</sup></b>	OUR EVALUATION

$$|q_p + q_{\bar{p}}|/e$$

A test of *CPT* invariance. Note that the comparison of the  $\bar{p}$  and  $p$  charge-to-mass ratios given above is much better determined. See also a similar test involving the electron.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;7 × 10<sup>-10</sup></b>	90	<sup>1</sup> HORI	11	SPEC $\bar{p}e^-$ -He atom
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2 × 10 <sup>-9</sup>	90	<sup>1</sup> HORI	06	SPEC $\bar{p}e^-$ -He atom
<1.0 × 10 <sup>-8</sup>	90	<sup>1</sup> HORI	03	SPEC $\bar{p}e^-$ <sup>4</sup> He, $\bar{p}e^-$ <sup>3</sup> He
<6 × 10 <sup>-8</sup>	90	<sup>1</sup> HORI	01	SPEC $\bar{p}e^-$ -He atom
<5 × 10 <sup>-7</sup>		<sup>2</sup> TORII	99	SPEC $\bar{p}e^-$ -He atom
<2 × 10 <sup>-5</sup>		<sup>3</sup> HUGHES	92	RVUE

<sup>1</sup> HORI 01, HORI 03, HORI 06, and HORI 11 use the more-precisely-known constraint on the  $\bar{p}$  charge-to-mass ratio of GABRIELSE 99 (see above) to get their results. Their results are not independent of the HORI 01, HORI 03, HORI 06, and HORI 11 values for  $|m_p - m_{\bar{p}}|/m_p$ , above.

<sup>2</sup> TORII 99 uses the more-precisely-known constraint on the  $\bar{p}$  charge-to-mass ratio of GABRIELSE 95 (see above) to get this result. This is not independent of the TORII 99 value for  $|m_p - m_{\bar{p}}|/m_p$ , above.

<sup>3</sup> HUGHES 92 uses recent measurements of Rydberg-energy and cyclotron-frequency ratios.

$$|q_p + q_e|/e$$

See BRESSI 11 for a summary of experiments on the neutrality of matter. See also “*n* CHARGE” in the neutron Listings.

VALUE	DOCUMENT ID	COMMENT
<b>&lt;1 × 10<sup>-21</sup></b>	<sup>1</sup> BRESSI	11 Neutrality of SF <sub>6</sub>
• • • We do not use the following data for averages, fits, limits, etc. • • •		
<3.2 × 10 <sup>-20</sup>	<sup>2</sup> SENGUPTA	00 binary pulsar
<0.8 × 10 <sup>-21</sup>	MARINELLI	84 Magnetic levitation
<1.0 × 10 <sup>-21</sup>	<sup>1</sup> DYLLA	73 Neutrality of SF <sub>6</sub>

<sup>1</sup> BRESSI 11 uses the method of DYLLA 73 but finds serious errors in that experiment that greatly reduce its accuracy. The BRESSI 11 limit assumes that  $n \rightarrow p e^- \nu_e$  conserves charge. Thus the limit applies equally to the charge of the neutron.

<sup>2</sup> SENGUPTA 00 uses the difference between the observed rate of rotational energy loss by the binary pulsar PSR B1913+16 and the rate predicted by general relativity to set this limit. See the paper for assumptions.

### p MAGNETIC MOMENT

See the “Note on Baryon Magnetic Moments” in the *A* Listings.

VALUE (μ <sub>N</sub> )	DOCUMENT ID	TECN	COMMENT
<b>2.7928473508 ± 0.0000000085</b>	MOHR 16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.792847356 ± 0.0000000023	MOHR 12	RVUE	2010 CODATA value
2.792847356 ± 0.0000000023	MOHR 08	RVUE	2006 CODATA value
2.792847351 ± 0.0000000028	MOHR 05	RVUE	2002 CODATA value
2.792847337 ± 0.0000000029	MOHR 99	RVUE	1998 CODATA value
2.792847386 ± 0.0000000063	COHEN 87	RVUE	1986 CODATA value
2.7928456 ± 0.0000011	COHEN 73	RVUE	1973 CODATA value

### $\bar{p}$ MAGNETIC MOMENT

A few early results have been omitted.

VALUE (μ <sub>N</sub> )	DOCUMENT ID	TECN	COMMENT
<b>−2.792845 ± 0.000012</b>	DISCIACCA 13	TRAP	Single $\bar{p}$ , Penning trap
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−2.7862 ± 0.0083	PASK 09	CNTR	$\bar{p}$ He <sup>+</sup> hyperfine structure
−2.8005 ± 0.0090	KREISSL 88	CNTR	$\bar{p}$ <sup>208</sup> Pb 11 → 10 X-ray
−2.817 ± 0.048	ROBERTS 78	CNTR	
−2.791 ± 0.021	HU 75	CNTR	Exotic atoms

$$(\mu_p + \mu_{\bar{p}})/\mu_p$$

A test of *CPT* invariance.

VALUE (units 10 <sup>−6</sup> )	DOCUMENT ID	TECN	COMMENT
<b>0 ± 5</b>	DISCIACCA 13	TRAP	Single $\bar{p}$ , Penning trap

### p ELECTRIC DIPOLE MOMENT

A nonzero value is forbidden by both *T* invariance and *P* invariance.

VALUE (10 <sup>−23</sup> ecm)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.54</b>	<sup>1</sup> DMITRIEV	03		Uses <sup>199</sup> Hg atom EDM

# Baryon Particle Listings

## p

- • • We do not use the following data for averages, fits, limits, etc. • • •
- 3.7 ± 6.3

< 400

130 ± 200

900 ±1400

700 ± 900

CHO

DZUBA

<sup>2</sup> WILKENING

<sup>3</sup> WILKENING

1G HARRISON

89

85

84

84

69

NMR

THEO Uses <sup>129</sup>Xe moment

MBR Molecular beam
- <sup>1</sup>DMITRIEV 03 calculates this limit from the limit on the electric dipole moment of the <sup>199</sup>Hg atom.
- <sup>2</sup>This WILKENING 84 value includes a finite-size effect and a magnetic effect.
- <sup>3</sup>This WILKENING 84 value is more cautious than the other and excludes the finite-size effect, which relies on uncertain nuclear integrals.

### p ELECTRIC POLARIZABILITY α<sub>p</sub>

For a very complete review of the “polarizability of the nucleon and Compton scattering,” see SCHUMACHER 05. His recommended values for the proton are α<sub>p</sub> = (12.0 ± 0.6) × 10<sup>−4</sup> fm<sup>3</sup> and β<sub>p</sub> = (1.9 ± 0.6) × 10<sup>−4</sup> fm<sup>3</sup>, almost exactly our averages.

VALUE (10 <sup>−4</sup> fm <sup>3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>11.2 ±0.4 OUR AVERAGE</b>			
10.65±0.35±0.36	MCGOVERN	13	RVUE χEFT + Compton scattering
12.1 ±1.1 ±0.5	<sup>1</sup> BEANE	03	EFT + γp
11.82±0.98± <sub>−0.98</sub> <sup>0.52</sup>	<sup>2</sup> BLANPIED	01	LEGS p( <i>γ</i> ,γ), p( <i>γ</i> ,π <sup>0</sup> ), p( <i>γ</i> ,π <sup>+</sup> )
11.9 ±0.5 ±1.3	<sup>3</sup> OLMOSDEL...	01	CNTR γp Compton scattering
12.1 ±0.8 ±0.5	<sup>4</sup> MACGIBBON	95	RVUE global average
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11.7 ±0.8 ±0.7	<sup>5</sup> BARANOV	01	RVUE Global average
12.5 ±0.6 ±0.9	MACGIBBON	95	CNTR γp Compton scattering
9.8 ±0.4 ±1.1	HALLIN	93	CNTR γp Compton scattering
10.62±1.25+1.07 −1.19−1.03	ZIEGER	92	CNTR γp Compton scattering
10.9 ±2.2 ±1.3	<sup>6</sup> FEDERSPIEL	91	CNTR γp Compton scattering
<sup>1</sup> BEANE 03 uses effective field theory and low-energy γp and γd Compton-scattering data. It also gets for the isoscalar polarizabilities (see the erratum) α <sub>N</sub> = (13.0 ± 1.9+3.9) −1.5 × 10 <sup>−4</sup> fm <sup>3</sup> and β <sub>N</sub> = (−1.8 ± 1.9±2.1) 0.9 × 10 <sup>−4</sup> fm <sup>3</sup> .			
<sup>2</sup> BLANPIED 01 gives α <sub>p</sub> + β <sub>p</sub> and α <sub>p</sub> − β <sub>p</sub> . The separate α <sub>p</sub> and β <sub>p</sub> are provided to us by A. Sandorfi. The first error above is statistics plus systematics; the second is from the model.			
<sup>3</sup> This OLMOSDELEON 01 result uses the TAPS data alone, and does not use the (re-evaluated) sum-rule constraint that α + β = (13.8 ± 0.4) × 10 <sup>−4</sup> fm <sup>3</sup> . See the paper for a discussion.			
<sup>4</sup> MACGIBBON 95 combine the results of ZIEGER 92, FEDERSPIEL 91, and their own experiment to get a “global average” in which model errors and systematic errors are treated in a consistent way. See MACGIBBON 95 for a discussion.			
<sup>5</sup> BARANOV 01 combines the results of 10 experiments from 1958 through 1995 to get a global average that takes into account both systematic and model errors and does not use the theoretical constraint on the sum α <sub>p</sub> + β <sub>p</sub> .			
<sup>6</sup> FEDERSPIEL 91 obtains for the (static) electric polarizability α <sub>p</sub> , defined in terms of the induced electric dipole moment by <b>D</b> = 4πε <sub>0</sub> α <sub>p</sub> <b>E</b> , the value (7.0 ± 2.2 ± 1.3) × 10 <sup>−4</sup> fm <sup>3</sup> .			

### p MAGNETIC POLARIZABILITY β<sub>p</sub>

The electric and magnetic polarizabilities are subject to a dispersion sum-rule constraint α + β = (14.2 ± 0.5) × 10<sup>−4</sup> fm<sup>3</sup>. Errors here are anticorrelated with those on α<sub>p</sub> due to this constraint.

VALUE (10 <sup>−4</sup> fm <sup>3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>2.5 ±0.4 OUR AVERAGE</b>			Error includes scale factor of 1.2.
3.15±0.35±0.36	MCGOVERN	13	RVUE χEFT + Compton scattering
3.4 ±1.1 ±0.1	<sup>1</sup> BEANE	03	EFT + γp
1.43±0.98± <sub>−0.98</sub> <sup>0.52</sup>	<sup>2</sup> BLANPIED	01	LEGS p( <i>γ</i> ,γ), p( <i>γ</i> ,π <sup>0</sup> ), p( <i>γ</i> ,π <sup>+</sup> )
1.2 ±0.7 ±0.5	<sup>3</sup> OLMOSDEL...	01	CNTR γp Compton scattering
2.1 ±0.8 ±0.5	<sup>4</sup> MACGIBBON	95	RVUE global average
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.3 ±0.9 ±0.7	<sup>5</sup> BARANOV	01	RVUE Global average
1.7 ±0.6 ±0.9	MACGIBBON	95	CNTR γp Compton scattering
4.4 ±0.4 ±1.1	HALLIN	93	CNTR γp Compton scattering
3.58±1.19+1.03 −1.25−1.07	ZIEGER	92	CNTR γp Compton scattering
3.3 ±2.2 ±1.3	FEDERSPIEL	91	CNTR γp Compton scattering
<sup>1</sup> BEANE 03 uses effective field theory and low-energy γp and γd Compton-scattering data. It also gets for the isoscalar polarizabilities (see the erratum) α <sub>N</sub> = (13.0 ± 1.9+3.9) −1.5 × 10 <sup>−4</sup> fm <sup>3</sup> and β <sub>N</sub> = (−1.8 ± 1.9±2.1) 0.9 × 10 <sup>−4</sup> fm <sup>3</sup> .			
<sup>2</sup> BLANPIED 01 gives α <sub>p</sub> + β <sub>p</sub> and α <sub>p</sub> − β <sub>p</sub> . The separate α <sub>p</sub> and β <sub>p</sub> are provided to us by A. Sandorfi. The first error above is statistics plus systematics; the second is from the model.			
<sup>3</sup> This OLMOSDELEON 01 result uses the TAPS data alone, and does not use the (re-evaluated) sum-rule constraint that α + β = (13.8 ± 0.4) × 10 <sup>−4</sup> fm <sup>3</sup> . See the paper for a discussion.			
<sup>4</sup> MACGIBBON 95 combine the results of ZIEGER 92, FEDERSPIEL 91, and their own experiment to get a “global average” in which model errors and systematic errors are treated in a consistent way. See MACGIBBON 95 for a discussion.			
<sup>5</sup> BARANOV 01 combines the results of 10 experiments from 1958 through 1995 to get a global average that takes into account both systematic and model errors and does not use the theoretical constraint on the sum α <sub>p</sub> + β <sub>p</sub> .			

### p CHARGE RADIUS

This is the rms electric charge radius, √⟨r<sub>E</sub><sup>2</sup>⟩.

Most measurements of the radius of the proton involve electron-proton interactions, and most of the more recent values agree with one another. The most precise of these is r<sub>p</sub> = 0.879(8) fm (BERNAUER 10). The CODATA 14 value (MOHR 16), obtained from the electronic results, is 0.8751(61). Compared to this CODATA value, however, a measurement using muonic hydrogen finds r<sub>p</sub> = 0.84087(39) fm (ANTOGNINI 13), which is 16 times more precise but differs by 5.6 standard deviations (using the CODATA 14 error).

Since POHL 10 (the first μp result), there has been a lot of discussion about the disagreement, especially concerning the modeling of muonic hydrogen. Here is an incomplete list of papers: DERUJULA 10, CLOET 11, DISTLER 11, DERUJULA 11, ARRINGTON 11, BERNAUER 11, HILL 11, LORENZ 14, KARSHENBOIM 14A, and PESET 15.

Until the difference between the ep and μp values is understood, it does not make sense to average the values together. For the present, we give both values. It is up to workers in this field to solve this puzzle.

See our 2014 edition (Chinese Physics C **38** 070001 (2014)) for values published before 2003.

VALUE (fm)	DOCUMENT ID	TECN	COMMENT
<b>0.8751 ±0.0061</b>	MOHR	16	RVUE 2014 CODATA value
<b>0.84087±0.00026±0.00029</b>	ANTOGNINI	13	LASR μp-atom Lamb shift
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.895 ±0.014 ±0.014	<sup>1</sup> LEE	15	SPEC Just 2010 Mainz data
0.916 ±0.024	LEE	15	SPEC World data, no Mainz
0.8775 ±0.0051	MOHR	12	RVUE 2010 CODATA, ep data
0.875 ±0.008 ±0.006	ZHAN	11	SPEC Recoil polarimetry
0.879 ±0.005 ±0.006	BERNAUER	10	SPEC ep → ep form factor
0.912 ±0.009 ±0.007	BORISYUK	10	reanalyzes old ep data
0.871 ±0.009 ±0.003	HILL	10	z-expansion reanalysis
0.84184±0.00036±0.00056	POHL	10	LASR See ANTOGNINI 13
0.8768 ±0.0069	MOHR	08	RVUE 2006 CODATA value
0.844 +0.008 −0.004	BELUSHKIN	07	Dispersion analysis
0.897 ±0.018	BLUNDEN	05	SICK 03 + 2γ correction
0.8750 ±0.0068	MOHR	05	RVUE 2002 CODATA value
0.895 ±0.010 ±0.013	SICK	03	ep → ep reanalysis
<sup>1</sup> Authors also provide values for combinations of all available data.			

### p MAGNETIC RADIUS

This is the rms magnetic radius, √⟨r<sub>M</sub><sup>2</sup>⟩.

VALUE (fm)	DOCUMENT ID	TECN	COMMENT
<b>0.776±0.034±0.017</b>	<sup>1</sup> LEE	15	SPEC Just 2010 Mainz data
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.914±0.035	LEE	15	SPEC World data, no Mainz
0.87 ±0.02	EPSTEIN	14	Using ep, en, ππ data
0.867±0.009±0.018	ZHAN	11	SPEC Recoil polarimetry
0.777±0.013±0.010	BERNAUER	10	SPEC ep → ep form factor
0.876±0.010±0.016	BORISYUK	10	Reanalyzes old ep → ep data
0.854±0.005	BELUSHKIN	07	Dispersion analysis
<sup>1</sup> Authors also provide values for a combination of all available data.			

### p MEAN LIFE

A test of baryon conservation. See the “p Partial Mean Lives” section below for limits for identified final states. The limits here are to “anything” or are for “disappearance” modes of a bound proton (p) or (n). See also the 3ν modes in the “Partial Mean Lives” section. Table 1 of BACK 03 is a nice summary.

LIMIT (years)	PARTICLE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;5.8 × 10<sup>29</sup></b>	<b>n</b>	90	<sup>1</sup> ARAKI	06	KLND n → invisible
<b>&gt;2.1 × 10<sup>29</sup></b>	<b>p</b>	90	<sup>2</sup> AHMED	04	SNO p → invisible
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>1.9 × 10 <sup>29</sup>	n	90	<sup>2</sup> AHMED	04	SNO n → invisible
>1.8 × 10 <sup>25</sup>	n	90	<sup>3</sup> BACK	03	BORX
>1.1 × 10 <sup>26</sup>	p	90	<sup>3</sup> BACK	03	BORX
>3.5 × 10 <sup>28</sup>	p	90	<sup>4</sup> ZDESENKO	03	p → invisible
>1 × 10 <sup>28</sup>	p	90	<sup>5</sup> AHMAD	02	SNO p → invisible
>4 × 10 <sup>23</sup>	p	95	TRETYAK	01	d → n + ?
>1.9 × 10 <sup>24</sup>	p	90	<sup>6</sup> BERNABEI	00B	DAMA
>1.6 × 10 <sup>25</sup>	p, n	7,8	EVANS	77	
>3 × 10 <sup>23</sup>	p	8	DIX	70	CNTR
>3 × 10 <sup>23</sup>	p, n	8,9	FLEROV	58	



See key on page 601

# Baryon Particle Listings

*p*

- <sup>1</sup> ARAKI 06 looks for signs of de-excitation of the residual nucleus after disappearance of a neutron from the  $s$  shell of  $^{12}\text{C}$ .
- <sup>2</sup> AHMED 04 looks for  $\gamma$  rays from the de-excitation of a residual  $^{15}\text{O}^*$  or  $^{15}\text{N}^*$  following the disappearance of a neutron or proton in  $^{16}\text{O}$ .
- <sup>3</sup> BACK 03 looks for decays of unstable nuclides left after  $N$  decays of parent  $^{12}\text{C}$ ,  $^{13}\text{C}$ ,  $^{16}\text{O}$  nuclei. These are "invisible channel" limits.
- <sup>4</sup> ZDESENKO 03 gets this limit on proton disappearance in deuterium by analyzing SNO data in AHMAD 02.
- <sup>5</sup> AHMAD 02 (see its footnote 7) looks for neutrons left behind after the disappearance of the proton in deuterons.
- <sup>6</sup> BERNABEI 00b looks for the decay of a  $^{128}_{53}\text{I}$  nucleus following the disappearance of a proton in the otherwise-stable  $^{129}_{54}\text{Xe}$  nucleus.
- <sup>7</sup> EVANS 77 looks for the daughter nuclide  $^{129}\text{Xe}$  from possible  $^{130}\text{Te}$  decays in ancient Te ore samples.
- <sup>8</sup> This mean-life limit has been obtained from a half-life limit by dividing the latter by  $\ln(2) = 0.693$ .
- <sup>9</sup> FLEROV 58 looks for the spontaneous fission of a  $^{232}\text{Th}$  nucleus after the disappearance of one of its nucleons.

## $\bar{p}$ MEAN LIFE

Of the two astrophysical limits here, that of GEER 00D involves considerably more refinements in its modeling. The other limits come from direct observations of stored antiprotons. See also " $\bar{p}$  Partial Mean Lives" after " $p$  Partial Mean Lives," below, for exclusive-mode limits. The best (lifetime/branching fraction) limit there is  $7 \times 10^5$  years, for  $\bar{p} \rightarrow e^- \gamma$ . We advance only the exclusive-mode limits to our Summary Tables.

LIMIT (years)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
>8 $\times 10^5$	90		<sup>1</sup> GEER 00D		$\bar{p}/p$ ratio, cosmic rays
>0.28			GABRIELSE 90	TRAP	Penning trap
>0.08	90	1	BELL 79	CNTR	Storage ring
>1 $\times 10^7$			GOLDEN 79	SPEC	$\bar{p}/p$ ratio, cosmic rays
>3.7 $\times 10^{-3}$			BREGMAN 78	CNTR	Storage ring
<sup>1</sup> GEER 00D uses agreement between a model of galactic $\bar{p}$ production and propagation and the observed $\bar{p}/p$ cosmic-ray spectrum to set this limit.					

## $p$ DECAY MODES

See the "Note on Nucleon Decay" in our 1994 edition (Phys. Rev. **D50**, 1173) for a short review.

The "partial mean life" limits tabulated here are the limits on  $\tau/B_j$ , where  $\tau$  is the total mean life and  $B_j$  is the branching fraction for the mode in question. For  $N$  decays,  $p$  and  $n$  indicate proton and neutron partial lifetimes.

Mode	Partial mean life ( $10^{30}$ years)	Confidence level
<b>Antilepton + meson</b>		
$\tau_1 \quad N \rightarrow e^+ \pi$	> 2000 ( $n$ ), > 8200 ( $p$ )	90%
$\tau_2 \quad N \rightarrow \mu^+ \pi$	> 1000 ( $n$ ), > 6600 ( $p$ )	90%
$\tau_3 \quad N \rightarrow \nu \pi$	> 1100 ( $n$ ), > 390 ( $p$ )	90%
$\tau_4 \quad p \rightarrow e^+ \eta$	> 4200	90%
$\tau_5 \quad p \rightarrow \mu^+ \eta$	> 1300	90%
$\tau_6 \quad n \rightarrow \nu \eta$	> 158	90%
$\tau_7 \quad N \rightarrow e^+ \rho$	> 217 ( $n$ ), > 710 ( $p$ )	90%
$\tau_8 \quad N \rightarrow \mu^+ \rho$	> 228 ( $n$ ), > 160 ( $p$ )	90%
$\tau_9 \quad N \rightarrow \nu \rho$	> 19 ( $n$ ), > 162 ( $p$ )	90%
$\tau_{10} \quad p \rightarrow e^+ \omega$	> 320	90%
$\tau_{11} \quad p \rightarrow \mu^+ \omega$	> 780	90%
$\tau_{12} \quad n \rightarrow \nu \omega$	> 108	90%
$\tau_{13} \quad N \rightarrow e^+ K$	> 17 ( $n$ ), > 1000 ( $p$ )	90%
$\tau_{14} \quad p \rightarrow e^+ K_S^0$		
$\tau_{15} \quad p \rightarrow e^+ K_L^0$		
$\tau_{16} \quad N \rightarrow \mu^+ K$	> 26 ( $n$ ), > 1600 ( $p$ )	90%
$\tau_{17} \quad p \rightarrow \mu^+ K_S^0$		
$\tau_{18} \quad p \rightarrow \mu^+ K_L^0$		
$\tau_{19} \quad N \rightarrow \nu K$	> 86 ( $n$ ), > 5900 ( $p$ )	90%
$\tau_{20} \quad n \rightarrow \nu K_S^0$	> 260	90%
$\tau_{21} \quad p \rightarrow e^+ K^*(892)^0$	> 84	90%
$\tau_{22} \quad N \rightarrow \nu K^*(892)$	> 78 ( $n$ ), > 51 ( $p$ )	90%

## Antilepton + mesons

Mode	Partial mean life (years)	Confidence level
$\tau_{23} \quad p \rightarrow e^+ \pi^+ \pi^-$	> 82	90%
$\tau_{24} \quad p \rightarrow e^+ \pi^0 \pi^0$	> 147	90%
$\tau_{25} \quad n \rightarrow e^+ \pi^- \pi^0$	> 52	90%
$\tau_{26} \quad p \rightarrow \mu^+ \pi^+ \pi^-$	> 133	90%
$\tau_{27} \quad p \rightarrow \mu^+ \pi^0 \pi^0$	> 101	90%
$\tau_{28} \quad n \rightarrow \mu^+ \pi^- \pi^0$	> 74	90%
$\tau_{29} \quad n \rightarrow e^+ K^0 \pi^-$	> 18	90%

## Lepton + meson

$\tau_{30} \quad n \rightarrow e^- \pi^+$	> 65	90%
$\tau_{31} \quad n \rightarrow \mu^- \pi^+$	> 49	90%
$\tau_{32} \quad n \rightarrow e^- \rho^+$	> 62	90%
$\tau_{33} \quad n \rightarrow \mu^- \rho^+$	> 7	90%
$\tau_{34} \quad n \rightarrow e^- K^+$	> 32	90%
$\tau_{35} \quad n \rightarrow \mu^- K^+$	> 57	90%

## Lepton + mesons

$\tau_{36} \quad p \rightarrow e^- \pi^+ \pi^+$	> 30	90%
$\tau_{37} \quad n \rightarrow e^- \pi^+ \pi^0$	> 29	90%
$\tau_{38} \quad p \rightarrow \mu^- \pi^+ \pi^+$	> 17	90%
$\tau_{39} \quad n \rightarrow \mu^- \pi^+ \pi^0$	> 34	90%
$\tau_{40} \quad p \rightarrow e^- \pi^+ K^+$	> 75	90%
$\tau_{41} \quad p \rightarrow \mu^- \pi^+ K^+$	> 245	90%

## Antilepton + photon(s)

$\tau_{42} \quad p \rightarrow e^+ \gamma$	> 670	90%
$\tau_{43} \quad p \rightarrow \mu^+ \gamma$	> 478	90%
$\tau_{44} \quad n \rightarrow \nu \gamma$	> 550	90%
$\tau_{45} \quad p \rightarrow e^+ \gamma \gamma$	> 100	90%
$\tau_{46} \quad n \rightarrow \nu \gamma \gamma$	> 219	90%

## Antilepton + single massless

$\tau_{47} \quad p \rightarrow e^+ X$	> 790	90%
$\tau_{48} \quad p \rightarrow \mu^+ X$	> 410	90%

## Three (or more) leptons

$\tau_{49} \quad p \rightarrow e^+ e^+ e^-$	> 793	90%
$\tau_{50} \quad p \rightarrow e^+ \mu^+ \mu^-$	> 359	90%
$\tau_{51} \quad p \rightarrow e^+ \nu \nu$	> 170	90%
$\tau_{52} \quad n \rightarrow e^+ e^- \nu$	> 257	90%
$\tau_{53} \quad n \rightarrow \mu^+ e^- \nu$	> 83	90%
$\tau_{54} \quad n \rightarrow \mu^+ \mu^- \nu$	> 79	90%
$\tau_{55} \quad p \rightarrow \mu^+ e^+ e^-$	> 529	90%
$\tau_{56} \quad p \rightarrow \mu^+ \mu^+ \mu^-$	> 675	90%
$\tau_{57} \quad p \rightarrow \mu^+ \nu \nu$	> 220	90%
$\tau_{58} \quad p \rightarrow e^- \mu^+ \mu^+$	> 6	90%
$\tau_{59} \quad n \rightarrow 3\nu$	> $5 \times 10^{-4}$	90%
$\tau_{60} \quad n \rightarrow 5\nu$		

## Inclusive modes

$\tau_{61} \quad N \rightarrow e^+ \text{anything}$	> 0.6 ( $n, p$ )	90%
$\tau_{62} \quad N \rightarrow \mu^+ \text{anything}$	> 12 ( $n, p$ )	90%
$\tau_{63} \quad N \rightarrow \nu \text{anything}$		
$\tau_{64} \quad N \rightarrow e^+ \pi^0 \text{anything}$	> 0.6 ( $n, p$ )	90%
$\tau_{65} \quad N \rightarrow 2 \text{ bodies, } \nu\text{-free}$		

## $\Delta B = 2$ dinucleon modes

The following are lifetime limits per iron nucleus.

$\tau_{66} \quad pp \rightarrow \pi^+ \pi^+$	> 72.2	90%
$\tau_{67} \quad pn \rightarrow \pi^+ \pi^0$	> 170	90%
$\tau_{68} \quad nn \rightarrow \pi^+ \pi^-$	> 0.7	90%
$\tau_{69} \quad nn \rightarrow \pi^0 \pi^0$	> 404	90%
$\tau_{70} \quad pp \rightarrow K^+ K^+$	> 170	90%
$\tau_{71} \quad pp \rightarrow e^+ e^+$	> 5.8	90%
$\tau_{72} \quad pp \rightarrow e^+ \mu^+$	> 3.6	90%
$\tau_{73} \quad pp \rightarrow \mu^+ \mu^+$	> 1.7	90%
$\tau_{74} \quad pn \rightarrow e^+ \bar{\nu}$	> 260	90%
$\tau_{75} \quad pn \rightarrow \mu^+ \bar{\nu}$	> 200	90%
$\tau_{76} \quad pn \rightarrow \tau^+ \bar{\nu}_\tau$	> 29	90%
$\tau_{77} \quad nn \rightarrow \nu_e \bar{\nu}_e$	> 1.4	90%
$\tau_{78} \quad nn \rightarrow \nu_\mu \bar{\nu}_\mu$	> 1.4	90%
$\tau_{79} \quad pn \rightarrow \text{invisible}$	> $2.1 \times 10^{-5}$	90%
$\tau_{80} \quad pp \rightarrow \text{invisible}$	> $5 \times 10^{-5}$	90%

## $\bar{p}$ DECAY MODES

Mode	Partial mean life (years)	Confidence level
$\tau_{81} \quad \bar{p} \rightarrow e^- \gamma$	> $7 \times 10^5$	90%
$\tau_{82} \quad \bar{p} \rightarrow \mu^- \gamma$	> $5 \times 10^4$	90%
$\tau_{83} \quad \bar{p} \rightarrow e^- \pi^0$	> $4 \times 10^5$	90%
$\tau_{84} \quad \bar{p} \rightarrow \mu^- \pi^0$	> $5 \times 10^4$	90%
$\tau_{85} \quad \bar{p} \rightarrow e^- \eta$	> $2 \times 10^4$	90%

# Baryon Particle Listings

$\rho$

$\tau_{86}$	$\bar{p} \rightarrow \mu^- \eta$	$> 8 \times 10^3$	90%
$\tau_{87}$	$\bar{p} \rightarrow e^- K_S^0$	$> 900$	90%
$\tau_{88}$	$\bar{p} \rightarrow \mu^- K_S^0$	$> 4 \times 10^3$	90%
$\tau_{89}$	$\bar{p} \rightarrow e^- K_L^0$	$> 9 \times 10^3$	90%
$\tau_{90}$	$\bar{p} \rightarrow \mu^- K_L^0$	$> 7 \times 10^3$	90%
$\tau_{91}$	$\bar{p} \rightarrow e^- \gamma \gamma$	$> 2 \times 10^4$	90%
$\tau_{92}$	$\bar{p} \rightarrow \mu^- \gamma \gamma$	$> 2 \times 10^4$	90%
$\tau_{93}$	$\bar{p} \rightarrow e^- \omega$	$> 200$	90%

## $\rho$ PARTIAL MEAN LIVES

The “partial mean life” limits tabulated here are the limits on  $\tau/B_i$ , where  $\tau$  is the total mean life for the proton and  $B_i$  is the branching fraction for the mode in question.

Decaying particle:  $p$  = proton,  $n$  = bound neutron. The same event may appear under more than one partial decay mode. Background estimates may be accurate to a factor of two.

## Antilepton + meson

### $\tau(N \rightarrow e^+ \pi)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt;2000</b>	<b><math>n</math></b>	<b>90</b>	<b>0</b>	<b>0.27</b>	NISHINO 12	SKAM
<b>&gt;8200</b>	<b><math>p</math></b>	<b>90</b>	<b>0</b>	<b>0.3</b>	NISHINO 09	SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

$> 540$	$p$	90	0	0.2	MCGREW 99	IMB3
$> 158$	$n$	90	3	5	MCGREW 99	IMB3
$>1600$	$p$	90	0	0.1	SHIOZAWA 98	SKAM
$> 70$	$p$	90	0	0.5	BERGER 91	FREJ
$> 70$	$n$	90	0	$\leq 0.1$	BERGER 91	FREJ
$> 550$	$p$	90	0	0.7	<sup>1</sup> BECKER-SZ... 90	IMB3
$> 260$	$p$	90	0	$<0.04$	HIRATA 89c	KAMI
$> 130$	$n$	90	0	$<0.2$	HIRATA 89c	KAMI
$> 310$	$p$	90	0	0.6	SEIDEL 88	IMB
$> 100$	$n$	90	0	1.6	SEIDEL 88	IMB
$> 1.3$	$n$	90	0		BARTELT 87	SLOUD
$> 1.3$	$p$	90	0		BARTELT 87	SLOUD
$> 250$	$p$	90	0	0.3	HAINES 86	IMB
$> 31$	$n$	90	8	9	HAINES 86	IMB
$> 64$	$p$	90	0	$<0.4$	ARISAKA 85	KAMI
$> 26$	$n$	90	0	$<0.7$	ARISAKA 85	KAMI
$> 82$	$p$ (free)	90	0	0.2	BLEWITT 85	IMB
$> 250$	$p$	90	0	0.2	BLEWITT 85	IMB
$> 25$	$n$	90	4	4	PARK 85	IMB
$> 15$	$p, n$	90	0		BATTISTONI 84	NUSX
$> 0.5$	$p$	90	1	0.3	<sup>2</sup> BARTELT 83	SLOUD
$> 0.5$	$n$	90	1	0.3	<sup>2</sup> BARTELT 83	SLOUD
$> 5.8$	$p$	90	2		<sup>3</sup> KRISHNA... 82	KOLR
$> 5.8$	$n$	90	2		<sup>3</sup> KRISHNA... 82	KOLR
$> 0.1$	$n$	90			<sup>4</sup> GURR 67	CNTR

<sup>1</sup> This BECKER-SZENDY 90 result includes data from SEIDEL 88.

<sup>2</sup> Limit based on zero events.

<sup>3</sup> We have calculated 90% CL limit from 1 confined event.

<sup>4</sup> We have converted half-life to 90% CL mean life.

### $\tau(N \rightarrow \mu^+ \pi)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt;1000</b>	<b><math>n</math></b>	<b>90</b>	<b>1</b>	<b>0.43</b>	NISHINO 12	SKAM
<b>&gt;6600</b>	<b><math>p</math></b>	<b>90</b>	<b>0</b>	<b>0.3</b>	NISHINO 09	SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

$> 473$	$p$	90	0	0.6	MCGREW 99	IMB3
$> 90$	$n$	90	1	1.9	MCGREW 99	IMB3
$> 81$	$p$	90	0	0.2	BERGER 91	FREJ
$> 35$	$n$	90	1	1.0	BERGER 91	FREJ
$> 230$	$p$	90	0	$<0.07$	HIRATA 89c	KAMI
$> 100$	$n$	90	0	$<0.2$	HIRATA 89c	KAMI
$> 270$	$p$	90	0	0.5	SEIDEL 88	IMB
$> 63$	$n$	90	0	0.5	SEIDEL 88	IMB
$> 76$	$p$	90	2	1	HAINES 86	IMB
$> 23$	$n$	90	8	7	HAINES 86	IMB
$> 46$	$p$	90	0	$<0.7$	ARISAKA 85	KAMI
$> 20$	$n$	90	0	$<0.4$	ARISAKA 85	KAMI
$> 59$	$p$ (free)	90	0	0.2	BLEWITT 85	IMB
$> 100$	$p$	90	1	0.4	BLEWITT 85	IMB
$> 38$	$n$	90	1	4	PARK 85	IMB
$> 10$	$p, n$	90	0		BATTISTONI 84	NUSX
$> 1.3$	$p, n$	90	0		ALEKSEEV 81	BAKS

### $\tau(N \rightarrow \nu \pi)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt; 390</b>	<b><math>p</math></b>	<b>90</b>	<b>52.8</b>		ABE 14E	SKAM
<b>&gt;1100</b>	<b><math>n</math></b>	<b>90</b>	<b>19.1</b>		ABE 14E	SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

$> 16$	$p$	90	6	6.7	WALL 00B	SOU2
$> 39$	$n$	90	4	3.8	WALL 00B	SOU2
$> 10$	$p$	90	15	20.3	MCGREW 99	IMB3
$> 112$	$n$	90	6	6.6	MCGREW 99	IMB3
$> 13$	$n$	90	1	1.2	BERGER 89	FREJ
$> 10$	$p$	90	11	14	BERGER 89	FREJ
$> 25$	$p$	90	32	32.8	<sup>1</sup> HIRATA 89c	KAMI
$> 100$	$n$	90	1	3	HIRATA 89c	KAMI
$> 6$	$n$	90	73	60	HAINES 86	IMB
$> 2$	$p$	90	16	13	KAJITA 86	KAMI
$> 40$	$n$	90	0	1	KAJITA 86	KAMI
$> 7$	$n$	90	28	19	PARK 85	IMB
$> 7$	$n$	90	0		BATTISTONI 84	NUSX
$> 2$	$p$	90	$\leq 3$		BATTISTONI 84	NUSX
$> 5.8$	$p$	90	1		<sup>2</sup> KRISHNA... 82	KOLR
$> 0.3$	$p$	90	2		<sup>3</sup> CHERRY 81	HOME
$> 0.1$	$p$	90			<sup>4</sup> GURR 67	CNTR

<sup>1</sup> In estimating the background, this HIRATA 89c limit (as opposed to the later limits of WALL 00B and MCGREW 99) does not take into account present understanding that the flux of  $\nu_\mu$  originating in the upper atmosphere is depleted. Doing so would reduce the background and thus also would reduce the limit here.

<sup>2</sup> We have calculated 90% CL limit from 1 confined event.

<sup>3</sup> We have converted 2 possible events to 90% CL limit.

<sup>4</sup> We have converted half-life to 90% CL mean life.

### $\tau(p \rightarrow e^+ \eta)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt;4200</b>	<b><math>p</math></b>	<b>90</b>	<b>0</b>	<b>0.44</b>	NISHINO 12	SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

$> 81$	$p$	90	1	1.7	WALL 00B	SOU2
$> 313$	$p$	90	0	0.2	MCGREW 99	IMB3
$> 44$	$p$	90	0	0.1	BERGER 91	FREJ
$> 140$	$p$	90	0	$<0.04$	HIRATA 89c	KAMI
$> 100$	$p$	90	0	0.6	SEIDEL 88	IMB
$> 200$	$p$	90	5	3.3	HAINES 86	IMB
$> 64$	$p$	90	0	$<0.8$	ARISAKA 85	KAMI
$> 64$	$p$ (free)	90	5	6.5	BLEWITT 85	IMB
$> 200$	$p$	90	5	4.7	BLEWITT 85	IMB
$> 1.2$	$p$	90	2		<sup>1</sup> CHERRY 81	HOME

<sup>1</sup> We have converted 2 possible events to 90% CL limit.

### $\tau(p \rightarrow \mu^+ \eta)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt;1300</b>	<b><math>p</math></b>	<b>90</b>	<b>2</b>	<b>0.49</b>	NISHINO 12	SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

$> 89$	$p$	90	0	1.6	WALL 00B	SOU2
$> 126$	$p$	90	3	2.8	MCGREW 99	IMB3
$> 26$	$p$	90	1	0.8	BERGER 91	FREJ
$> 69$	$p$	90	1	$<0.08$	HIRATA 89c	KAMI
$> 1.3$	$p$	90	0	0.7	PHILLIPS 89	HPW
$> 34$	$p$	90	1	1.5	SEIDEL 88	IMB
$> 46$	$p$	90	7	6	HAINES 86	IMB
$> 26$	$p$	90	1	$<0.8$	ARISAKA 85	KAMI
$> 17$	$p$ (free)	90	6	6	BLEWITT 85	IMB
$> 46$	$p$	90	7	8	BLEWITT 85	IMB

### $\tau(n \rightarrow \nu \eta)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt;158</b>	<b><math>n</math></b>	<b>90</b>	<b>0</b>	<b>1.2</b>	MCGREW 99	IMB3

• • • We do not use the following data for averages, fits, limits, etc. • • •

$> 71$	$n$	90	2	3.7	WALL 00B	SOU2
$> 29$	$n$	90	0	0.9	BERGER 89	FREJ
$> 54$	$n$	90	2	0.9	HIRATA 89c	KAMI
$> 16$	$n$	90	3	2.1	SEIDEL 88	IMB
$> 25$	$n$	90	7	6	HAINES 86	IMB
$> 30$	$n$	90	0	0.4	KAJITA 86	KAMI
$> 18$	$n$	90	4	3	PARK 85	IMB
$> 0.6$	$n$	90	2		<sup>1</sup> CHERRY 81	HOME

<sup>1</sup> We have converted 2 possible events to 90% CL limit.

### $\tau(N \rightarrow e^+ \rho)$

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
<b>&gt;710</b>	<b><math>p</math></b>	<b>90</b>	<b>0</b>	<b>0.35</b>	NISHINO 12	SKAM
<b>&gt;217</b>	<b><math>n</math></b>	<b>90</b>	<b>4</b>	<b>4.8</b>	MCGREW 99	IMB3

See key on page 601

## Baryon Particle Listings

 $p$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 70	$n$	90	1	0.38	NISHINO	12	SKAM
> 29	$p$	90	0	2.2	BERGER	91	FREJ
> 41	$n$	90	0	1.4	BERGER	91	FREJ
> 75	$p$	90	2	2.7	HIRATA	89c	KAMI
> 58	$n$	90	0	1.9	HIRATA	89c	KAMI
> 38	$n$	90	2	4.1	SEIDEL	88	IMB
> 1.2	$p$	90	0		BARTELT	87	SOU
> 1.5	$n$	90	0		BARTELT	87	SOU
> 17	$p$	90	7	7	HAINES	86	IMB
> 14	$n$	90	9	4	HAINES	86	IMB
> 12	$p$	90	0	<1.2	ARISAKA	85	KAMI
> 6	$n$	90	2	<1	ARISAKA	85	KAMI
> 6.7	$p$ (free)	90	6	6	BLEWITT	85	IMB
> 17	$p$	90	7	7	BLEWITT	85	IMB
> 12	$n$	90	4	2	PARK	85	IMB
> 0.6	$n$	90	1	0.3	<sup>1</sup> BARTELT	83	SOU
> 0.5	$p$	90	1	0.3	<sup>1</sup> BARTELT	83	SOU
> 9.8	$p$	90	1		<sup>2</sup> KRISHNA...	82	KOLR
> 0.8	$p$	90	2		<sup>3</sup> CHERRY	81	HOME

<sup>1</sup> Limit based on zero events.<sup>2</sup> We have calculated 90% CL limit from 0 confined events.<sup>3</sup> We have converted 2 possible events to 90% CL limit. $\tau(N \rightarrow \mu^+ \rho)$  $\tau_8$ 

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>160	$p$	90	1	0.42	NISHINO	12 SKAM
>228	$n$	90	3	9.5	MCGREW	99 IMB3

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 36	$n$	90	0	0.29	NISHINO	12	SKAM
> 12	$p$	90	0	0.5	BERGER	91	FREJ
> 22	$n$	90	0	1.1	BERGER	91	FREJ
>110	$p$	90	0	1.7	HIRATA	89c	KAMI
> 23	$n$	90	1	1.8	HIRATA	89c	KAMI
> 4.3	$p$	90	0	0.7	PHILLIPS	89	HPW
> 30	$p$	90	0	0.5	SEIDEL	88	IMB
> 11	$n$	90	1	1.1	SEIDEL	88	IMB
> 16	$p$	90	4	4.5	HAINES	86	IMB
> 7	$n$	90	6	5	HAINES	86	IMB
> 12	$p$	90	0	<0.7	ARISAKA	85	KAMI
> 5	$n$	90	1	<1.2	ARISAKA	85	KAMI
> 5.5	$p$ (free)	90	4	5	BLEWITT	85	IMB
> 16	$p$	90	4	5	BLEWITT	85	IMB
> 9	$n$	90	1	2	PARK	85	IMB

 $\tau(N \rightarrow \nu \rho)$  $\tau_9$ 

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>162	$p$	90	18	21.7	MCGREW	99 IMB3
> 19	$n$	90	0	0.5	SEIDEL	88 IMB

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 9	$n$	90	4	2.4	BERGER	89	FREJ
> 24	$p$	90	0	0.9	BERGER	89	FREJ
> 27	$p$	90	5	1.5	HIRATA	89c	KAMI
> 13	$n$	90	4	3.6	HIRATA	89c	KAMI
> 13	$p$	90	1	1.1	SEIDEL	88	IMB
> 8	$p$	90	6	5	HAINES	86	IMB
> 2	$n$	90	15	10	HAINES	86	IMB
> 11	$p$	90	2	1	KAJITA	86	KAMI
> 4	$n$	90	2	2	KAJITA	86	KAMI
> 4.1	$p$ (free)	90	6	7	BLEWITT	85	IMB
> 8.4	$p$	90	6	5	BLEWITT	85	IMB
> 2	$n$	90	7	3	PARK	85	IMB
> 0.9	$p$	90	2		<sup>1</sup> CHERRY	81	HOME
> 0.6	$n$	90	2		<sup>1</sup> CHERRY	81	HOME

<sup>1</sup> We have converted 2 possible events to 90% CL limit. $\tau(p \rightarrow e^+ \omega)$  $\tau_{10}$ 

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>320	$p$	90	1	0.53	NISHINO	12 SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

>107	$p$	90	7	10.8	MCGREW	99	IMB3
> 17	$p$	90	0	1.1	BERGER	91	FREJ
> 45	$p$	90	2	1.45	HIRATA	89c	KAMI
> 26	$p$	90	1	1.0	SEIDEL	88	IMB
> 1.5	$p$	90	0		BARTELT	87	SOU
> 37	$p$	90	6	5.3	HAINES	86	IMB
> 25	$p$	90	1	<1.4	ARISAKA	85	KAMI
> 12	$p$ (free)	90	6	7.5	BLEWITT	85	IMB
> 37	$p$	90	6	5.7	BLEWITT	85	IMB
> 0.6	$p$	90	1	0.3	<sup>1</sup> BARTELT	83	SOU
> 9.8	$p$	90	1		<sup>2</sup> KRISHNA...	82	KOLR
> 2.8	$p$	90	2		<sup>3</sup> CHERRY	81	HOME

<sup>1</sup> Limit based on zero events.<sup>2</sup> We have calculated 90% CL limit from 0 confined events.<sup>3</sup> We have converted 2 possible events to 90% CL limit. $\tau(p \rightarrow \mu^+ \omega)$  $\tau_{11}$ 

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>780	$p$	90	0	0.48	NISHINO	12 SKAM

• • • We do not use the following data for averages, fits, limits, etc. • • •

>117	$p$	90	11	12.1	MCGREW	99	IMB3
> 11	$p$	90	0	1.0	BERGER	91	FREJ
> 57	$p$	90	2	1.9	HIRATA	89c	KAMI
> 4.4	$p$	90	0	0.7	PHILLIPS	89	HPW
> 10	$p$	90	2	1.3	SEIDEL	88	IMB
> 23	$p$	90	2	1	HAINES	86	IMB
> 6.5	$p$ (free)	90	9	8.7	BLEWITT	85	IMB
> 23	$p$	90	8	7	BLEWITT	85	IMB

 $\tau(n \rightarrow \nu \omega)$  $\tau_{12}$ 

<i>LIMIT</i> ( $10^{30}$ years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
>108	<i>n</i>	90	12	22.5	MCGREW 99	IMB3

• • • We do not use the following data for averages, fits, limits, etc. • • •

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 17	$n$	90	1	0.7	BERGER	89	FREJ
> 43	$n$	90	3	2.7	HIRATA	89c	KAMI
> 6	$n$	90	2	1.3	SEIDEL	88	IMB
> 12	$n$	90	6	6	HAINES	86	IMB
> 18	$n$	90	2	2	KAJITA	86	KAMI
> 16	$n$	90	1	2	PARK	85	IMB
> 2.0	$n$	90	2		<sup>1</sup> CHERRY	81	HOME

<sup>1</sup> We have converted 2 possible events to 90% CL limit. $\tau(N \rightarrow e^+ K)$  $\tau_{13}$ 

<u>LIMIT</u> <u>(10<sup>30</sup> years)</u>	<u>PARTICLE</u>	<u>CL%</u>	<u>EVTs</u>	<u>BKGD EST</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
>1000	p	90	6	4.7	KOBAYASHI	05 SKAM
> 17	n	90	35	29.4	MCGREW	99 IMB3

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 85	$p$	90	3	4.9	WALL	00	SOU2
> 31	$p$	90	23	25.2	MCGREW	99	IMB3
> 60	$p$	90	0		BERGER	91	FREJ
> 150	$p$	90	0	<0.27	HIRATA	89c	KAMI
> 70	$p$	90	0	1.8	SEIDEL	88	IMB
> 77	$p$	90	5	4.5	HAINES	86	IMB
> 38	$p$	90	0	<0.8	ARISAKA	85	KAMI
> 24	$p$ (free)	90	7	8.5	BLEWITT	85	IMB
> 77	$p$	90	5	4	BLEWITT	85	IMB
> 1.3	$p$	90	0		ALEKSEEV	81	BAKS
> 1.3	$n$	90	0		ALEKSEEV	81	BAKS

 $\tau(p \rightarrow e^+ K_S^0)$  $\tau_{14}$ 

<u>LIMIT</u> ( $10^{30}$ years)	<u>PARTICLE</u>	<u>CL%</u>	<u>EVTs</u>	<u>BKGD EST</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
>120	p	90	1	1.3	WALL	00
> 76	p	90	0	0.5	BERGER	91

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\tau(p \rightarrow e^+ K_L^0)$  $\tau_{15}$ 

<u>LIMIT</u> ( $10^{30}$ years)	<u>PARTICLE</u>	<u>CL%</u>	<u>EVTs</u>	<u>BKGD EST</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
• • •	We do not use the following data for averages, fits, limits, etc. • • •					
> 51	p	90	2	3.5	WALL	00
> 44	p	90	0	$\leq 0.1$	BERGER	91

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $\tau(N \rightarrow \mu^+ K)$  $\tau_{16}$ 

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>1600	<i>p</i>	90	13	13.2	REGIS	12 SKAM
> 26	<i>n</i>	90	20	28.4	MCGREW	99 IMB3

• • • We do not use the following data for averages, fits, limits, etc. • • •

>1300	$p$	90	3	3.9	KOBAYASHI	05	SKAM
> 120	$p$	90	0	<1.2	WALL	00	SOU2
> 120	$p$	90	4	7.2	MCGREW	99	IMB3
> 54	$p$	90	0		BERGER	91	FREJ
> 120	$p$	90	1	0.4	HIRATA	89c	KAMI
> 3.0	$p$	90	0	0.7	PHILLIPS	89	HPW
> 19	$p$	90	3	2.5	SEIDEL	88	IMB
> 1.5	$p$	90	0		<sup>1</sup> BARTELT	87	SOU
> 1.1	$n$	90	0		BARTELT	87	SOU
> 40	$p$	90	7	6	HAINES	86	IMB
> 19	$p$	90	1	<1.1	ARISAKA	85	KAMI
> 6.7	$p$ (free)	90	11	13	BLEWITT	85	IMB
> 40	$p$	90	7	8	BLEWITT	85	IMB
> 6	$p$	90	1		BATTISTONI	84	NUSX
> 0.6	$p$	90	0		<sup>2</sup> BARTELT	83	SOU
> 0.4	$n$	90	0		<sup>2</sup> BARTELT	83	SOU
> 5.8	$p$	90	2		<sup>3</sup> KRISHNA...	82	KOLR
> 2.0	$p$	90	0		CHERRY	81	HOME
> 0.2	$n$	90			<sup>4</sup> GURR	67	CNTR

# Baryon Particle Listings

$p$

<sup>1</sup> BARTELT 87 limit applies to  $p \rightarrow \mu^+ K_S^0$ .

<sup>2</sup> Limit based on zero events.

<sup>3</sup> We have calculated 90% CL limit from 1 confined event.

<sup>4</sup> We have converted half-life to 90% CL mean life.

## $\tau(p \rightarrow \mu^+ K_S^0)$

$\tau_{17}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>150	$p$	90	0	<0.8	WALL 00	SOU2
> 64	$p$	90	0	1.2	BERGER 91	FREJ

## $\tau(p \rightarrow \mu^+ K_L^0)$

$\tau_{18}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>83	$p$	90	0	0.4	WALL 00	SOU2
>44	$p$	90	0	$\leq 0.1$	BERGER 91	FREJ

## $\tau(N \rightarrow \nu K)$

$\tau_{19}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>5900</b>	$p$	<b>90</b>	<b>0</b>	<b>1.0</b>	ABE 14g	SKAM
> <b>86</b>	$n$	<b>90</b>	<b>0</b>	<b>2.4</b>	HIRATA 89c	KAMI
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 540	$p$	90	0	0.9	ASAKURA 15	KLND
>2300	$p$	90	0	1.3	KOBAYASHI 05	SKAM
> 26	$n$	90	16	9.1	WALL 00	SOU2
> 670	$p$	90			HAYATO 99	SKAM
> 151	$p$	90	15	21.4	MCGREW 99	IMB3
> 30	$n$	90	34	34.1	MCGREW 99	IMB3
> 43	$p$	90	1	1.54	<sup>1</sup> ALLISON 98	SOU2
> 15	$n$	90	1	1.8	BERGER 89	FREJ
> 15	$p$	90	1	1.8	BERGER 89	FREJ
> 100	$p$	90	9	7.3	HIRATA 89c	KAMI
> 0.28	$p$	90	0	0.7	PHILLIPS 89	HPW
> 0.3	$p$	90	0		BARTELT 87	SOD
> 0.75	$n$	90	0		<sup>2</sup> BARTELT 87	SOD
> 10	$p$	90	6	5	HAINES 86	IMB
> 15	$n$	90	3	5	HAINES 86	IMB
> 28	$p$	90	3	3	KAJITA 86	KAMI
> 32	$n$	90	0	1.4	KAJITA 86	KAMI
> 1.8	$p$ (free)	90	6	11	BLEWITT 85	IMB
> 9.6	$p$	90	6	5	BLEWITT 85	IMB
> 10	$n$	90	2	2	PARK 85	IMB
> 5	$n$	90	0		BATTISTONI 84	NUSX
> 2	$p$	90	0		BATTISTONI 84	NUSX
> 0.3	$n$	90	0		<sup>3</sup> BARTELT 83	SOD
> 0.1	$p$	90	0		<sup>3</sup> BARTELT 83	SOD
> 5.8	$p$	90	1		<sup>4</sup> KRISHNA... 82	KOLR
> 0.3	$n$	90	2		<sup>5</sup> CHERRY 81	HOME

<sup>1</sup> This ALLISON 98 limit is with no background subtraction; with subtraction the limit becomes  $> 46 \times 10^{30}$  years.

<sup>2</sup> BARTELT 87 limit applies to  $n \rightarrow \nu K_S^0$ .

<sup>3</sup> Limit based on zero events.

<sup>4</sup> We have calculated 90% CL limit from 1 confined event.

<sup>5</sup> We have converted 2 possible events to 90% CL limit.

## $\tau(n \rightarrow \nu K_S^0)$

$\tau_{20}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>260</b>	$n$	<b>90</b>	<b>34</b>	<b>30</b>	<sup>1</sup> KOBAYASHI 05	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 51	$n$	90	16	9.1	WALL 00	SOU2

<sup>1</sup> We have doubled the  $n \rightarrow \nu K^0$  limit given in KOBAYASHI 05 to obtain this  $n \rightarrow \nu K_S^0$  limit.

## $\tau(p \rightarrow e^+ K^*(892)^0)$

$\tau_{21}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>84</b>	$p$	<b>90</b>	<b>38</b>	<b>52.0</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>10	$p$	90	0	0.8	BERGER 91	FREJ
>52	$p$	90	2	1.55	HIRATA 89c	KAMI
>10	$p$	90	1	<1	ARISAKA 85	KAMI

## $\tau(N \rightarrow \nu K^*(892))$

$\tau_{22}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>51</b>	$p$	<b>90</b>	<b>7</b>	<b>9.1</b>	MCGREW 99	IMB3
> <b>78</b>	$n$	<b>90</b>	<b>40</b>	<b>50</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>22	$n$	90	0	2.1	BERGER 89	FREJ
>17	$p$	90	0	2.4	BERGER 89	FREJ
>20	$p$	90	5	2.1	HIRATA 89c	KAMI
>21	$n$	90	4	2.4	HIRATA 89c	KAMI
>10	$p$	90	7	6	HAINES 86	IMB
> 5	$n$	90	8	7	HAINES 86	IMB
> 8	$p$	90	3	2	KAJITA 86	KAMI
> 6	$n$	90	2	1.6	KAJITA 86	KAMI
> 5.8	$p$ (free)	90	10	16	BLEWITT 85	IMB
> 9.6	$p$	90	7	6	BLEWITT 85	IMB
> 7	$n$	90	1	4	PARK 85	IMB
> 2.1	$p$	90	1		<sup>1</sup> BATTISTONI 82	NUSX

<sup>1</sup> We have converted 1 possible event to 90% CL limit.

### Antilepton + mesons

## $\tau(p \rightarrow e^+ \pi^+ \pi^-)$

$\tau_{23}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>82</b>	$p$	<b>90</b>	<b>16</b>	<b>23.1</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>21	$p$	90	0	2.2	BERGER 91	FREJ

## $\tau(p \rightarrow e^+ \pi^0 \pi^0)$

$\tau_{24}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>147</b>	$p$	<b>90</b>	<b>2</b>	<b>0.8</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 38	$p$	90	1	0.5	BERGER 91	FREJ

## $\tau(n \rightarrow e^+ \pi^- \pi^0)$

$\tau_{25}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>52</b>	$n$	<b>90</b>	<b>38</b>	<b>34.2</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>32	$n$	90	1	0.8	BERGER 91	FREJ

## $\tau(p \rightarrow \mu^+ \pi^+ \pi^-)$

$\tau_{26}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>133</b>	$p$	<b>90</b>	<b>25</b>	<b>38.0</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 17	$p$	90	1	2.6	BERGER 91	FREJ
> 3.3	$p$	90	0	0.7	PHILLIPS 89	HPW

## $\tau(p \rightarrow \mu^+ \pi^0 \pi^0)$

$\tau_{27}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>101</b>	$p$	<b>90</b>	<b>3</b>	<b>1.6</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 33	$p$	90	1	0.9	BERGER 91	FREJ

## $\tau(n \rightarrow \mu^+ \pi^- \pi^0)$

$\tau_{28}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>74</b>	$n$	<b>90</b>	<b>17</b>	<b>20.8</b>	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>33	$n$	90	0	1.1	BERGER 91	FREJ

## $\tau(n \rightarrow e^+ K^0 \pi^-)$

$\tau_{29}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>18</b>	$n$	<b>90</b>	<b>1</b>	<b>0.2</b>	BERGER 91	FREJ

### Lepton + meson

## $\tau(n \rightarrow e^- \pi^+)$

$\tau_{30}$

<i>LIMIT</i> (10 <sup>30</sup> years)	<i>PARTICLE</i>	<i>CL%</i>	<i>EVTs</i>	<i>BKGD EST</i>	<i>DOCUMENT ID</i>	<i>TECN</i>
> <b>65</b>	$n$	<b>90</b>	<b>0</b>	<b>1.6</b>	SEIDEL 88	IMB
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>55	$n$	90	0	1.09	BERGER 91b	FREJ
>16	$n$	90	9	7	HAINES 86	IMB
>25	$n$	90	2	4	PARK 85	IMB

See key on page 601

## Baryon Particle Listings

 $p$  $\tau(n \rightarrow \mu^- \pi^+)$  731

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>49	$n$	90	0	0.5	SEIDEL 88	IMB
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>33	$n$	90	0	1.40	BERGER 91B	FREJ
> 2.7	$n$	90	0	0.7	PHILLIPS 89	HPW
>25	$n$	90	7	6	HAINES 86	IMB
>27	$n$	90	2	3	PARK 85	IMB

 $\tau(n \rightarrow e^- \rho^+)$  732

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>62	$n$	90	2	4.1	SEIDEL 88	IMB
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>12	$n$	90	13	6	HAINES 86	IMB
>12	$n$	90	5	3	PARK 85	IMB

 $\tau(n \rightarrow \mu^- \rho^+)$  733

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>7	$n$	90	1	1.1	SEIDEL 88	IMB
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>2.6	$n$	90	0	0.7	PHILLIPS 89	HPW
>9	$n$	90	7	5	HAINES 86	IMB
>9	$n$	90	2	2	PARK 85	IMB

 $\tau(n \rightarrow e^- K^+)$  734

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>32	$n$	90	3	2.96	BERGER 91B	FREJ
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 0.23	$n$	90	0	0.7	PHILLIPS 89	HPW

 $\tau(n \rightarrow \mu^- K^+)$  735

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>57	$n$	90	0	2.18	BERGER 91B	FREJ
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 4.7	$n$	90	0	0.7	PHILLIPS 89	HPW

## Lepton + mesons

 $\tau(p \rightarrow e^- \pi^+ \pi^+)$  736

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>30	$p$	90	1	2.50	BERGER 91B	FREJ
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 2.0	$p$	90	0	0.7	PHILLIPS 89	HPW

 $\tau(n \rightarrow e^- \pi^+ \pi^0)$  737

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>29	$n$	90	1	0.78	BERGER 91B	FREJ

 $\tau(p \rightarrow \mu^- \pi^+ \pi^+)$  738

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>17	$p$	90	1	1.72	BERGER 91B	FREJ
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 7.8	$p$	90	0	0.7	PHILLIPS 89	HPW

 $\tau(n \rightarrow \mu^- \pi^+ \pi^0)$  739

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>34	$n$	90	0	0.78	BERGER 91B	FREJ

 $\tau(p \rightarrow e^- \pi^+ K^+)$  740

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>75	$p$	90	81	127.2	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>20	$p$	90	3	2.50	BERGER 91B	FREJ

 $\tau(p \rightarrow \mu^- \pi^+ K^+)$  741

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>245	$p$	90	3	4.0	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 5	$p$	90	2	0.78	BERGER 91B	FREJ

## Antilepton + photon(s)

 $\tau(p \rightarrow e^+ \gamma)$  742

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>670	$p$	90	0	0.1	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>133	$p$	90	0	0.3	BERGER 91	FREJ
>460	$p$	90	0	0.6	SEIDEL 88	IMB
>360	$p$	90	0	0.3	HAINES 86	IMB
> 87	$p$ (free)	90	0	0.2	BLEWITT 85	IMB
>360	$p$	90	0	0.2	BLEWITT 85	IMB
> 0.1	$p$	90			<sup>1</sup> GURR 67	CNTR

<sup>1</sup> We have converted half-life to 90% CL mean life. $\tau(p \rightarrow \mu^+ \gamma)$  743

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>478	$p$	90	0	0.1	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>155	$p$	90	0	0.1	BERGER 91	FREJ
>380	$p$	90	0	0.5	SEIDEL 88	IMB
> 97	$p$	90	3	2	HAINES 86	IMB
> 61	$p$ (free)	90	0	0.2	BLEWITT 85	IMB
>280	$p$	90	0	0.6	BLEWITT 85	IMB
> 0.3	$p$	90			<sup>1</sup> GURR 67	CNTR

<sup>1</sup> We have converted half-life to 90% CL mean life. $\tau(n \rightarrow \nu \gamma)$  744

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>550	$n$	90			TAKHISTOV 15	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 28	$n$	90	163	144.7	MCGREW 99	IMB3
> 24	$n$	90	10	6.86	BERGER 91B	FREJ
> 9	$n$	90	73	60	HAINES 86	IMB
> 11	$n$	90	28	19	PARK 85	IMB

 $\tau(p \rightarrow e^+ \gamma \gamma)$  745

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>100	$p$	90	1	0.8	BERGER 91	FREJ

 $\tau(n \rightarrow \nu \gamma \gamma)$  746

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>219	$n$	90	5	7.5	MCGREW 99	IMB3

## Antilepton + single massless

 $\tau(p \rightarrow e^+ X)$  747

VALUE ( $10^{30}$ years)	CL%	DOCUMENT ID	TECN
>790	90	TAKHISTOV 15	SKAM

 $\tau(p \rightarrow \mu^+ X)$  748

VALUE ( $10^{30}$ years)	CL%	DOCUMENT ID	TECN
>410	90	TAKHISTOV 15	SKAM

## Three (or more) leptons

 $\tau(p \rightarrow e^+ e^+ e^-)$  749

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>793	$p$	90	0	0.5	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>147	$p$	90	0	0.1	BERGER 91	FREJ
>510	$p$	90	0	0.3	HAINES 86	IMB
> 89	$p$ (free)	90	0	0.5	BLEWITT 85	IMB
>510	$p$	90	0	0.7	BLEWITT 85	IMB

 $\tau(p \rightarrow e^+ \mu^+ \mu^-)$  750

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>359	$p$	90	1	0.9	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 81	$p$	90	0	0.16	BERGER 91	FREJ
> 5.0	$p$	90	0	0.7	PHILLIPS 89	HPW

 $\tau(p \rightarrow e^+ \nu \nu)$  751

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>170	$p$	90			<sup>1</sup> TAKHISTOV 14	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 17	$p$	90	152	153.7	MCGREW 99	IMB3
> 11	$p$	90	11	6.08	BERGER 91B	FREJ

<sup>1</sup> Allowed events at 90% CL are 459.

# Baryon Particle Listings

$p$

$\tau(n \rightarrow e^+ e^- \nu)$ <span>752</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>257	$n$	90	5	7.5	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 74	$n$	90	0	< 0.1	BERGER 91B	FREJ
> 45	$n$	90	5	5	HAINES 86	IMB
> 26	$n$	90	4	3	PARK 85	IMB

$\tau(n \rightarrow \mu^+ e^- \nu)$ <span>753</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>83	$n$	90	25	29.4	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>47	$n$	90	0	< 0.1	BERGER 91B	FREJ

$\tau(n \rightarrow \mu^+ \mu^- \nu)$ <span>754</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>79	$n$	90	100	145	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>42	$n$	90	0	1.4	BERGER 91B	FREJ
> 5.1	$n$	90	0	0.7	PHILLIPS 89	HPW
>16	$n$	90	14	7	HAINES 86	IMB
>19	$n$	90	4	7	PARK 85	IMB

$\tau(p \rightarrow \mu^+ e^- \nu)$ <span>755</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>529	$p$	90	0	1.0	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 91	$p$	90	0	$\leq 0.1$	BERGER 91	FREJ

$\tau(p \rightarrow \mu^+ \mu^- \nu)$ <span>756</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>675	$p$	90	0	0.3	MCGREW 99	IMB3
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>119	$p$	90	0	0.2	BERGER 91	FREJ
> 10.5	$p$	90	0	0.7	PHILLIPS 89	HPW
>190	$p$	90	1	0.1	HAINES 86	IMB
> 44	$p$ (free)	90	1	0.7	BLEWITT 85	IMB
>190	$p$	90	1	0.9	BLEWITT 85	IMB
> 2.1	$p$	90	1		1 BATTISTONI 82	NUSX

<sup>1</sup> We have converted 1 possible event to 90% CL limit.

$\tau(p \rightarrow \mu^+ \nu \nu)$ <span>757</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>220	$p$	90			1 TAKHISTOV 14	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 21	$p$	90	7	11.23	BERGER 91B	FREJ

<sup>1</sup> Allowed events at 90% CL are 286.

$\tau(p \rightarrow e^- \mu^+ \nu)$ <span>758</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>6.0	$p$	90	0	0.7	PHILLIPS 89	HPW

$\tau(n \rightarrow 3\nu)$ <span>759</span>						
See also the “to anything” and “disappearance” limits for bound nucleons in the “p Mean Life” data block just in front of the list of possible $p$ decay modes. Such modes could of course be to three (or five) neutrinos, and the limits are stronger, but we do not repeat them here.						

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>0.00049	$n$	90	2	2	1 SUZUKI 93B	KAMI
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>0.0023	$n$	90			2 GLICENSTEIN 97	KAMI
>0.00003	$n$	90	11	6.1	3 BERGER 91B	FREJ
>0.00012	$n$	90	7	11.2	3 BERGER 91B	FREJ
>0.0005	$n$	90	0		LEARNED 79	RVUE

<sup>1</sup> The SUZUKI 93B limit applies to any of  $\nu_e \nu_e \bar{\nu}_e$ ,  $\nu_\mu \nu_\mu \bar{\nu}_\mu$ , or  $\nu_\tau \nu_\tau \bar{\nu}_\tau$ .

<sup>2</sup> GLICENSTEIN 97 uses Kamioka data and the idea that the disappearance of the neutron’s magnetic moment should produce radiation.

<sup>3</sup> The first BERGER 91B limit is for  $n \rightarrow \nu_e \nu_e \bar{\nu}_e$ , the second is for  $n \rightarrow \nu_\mu \nu_\mu \bar{\nu}_\mu$ .

$\tau(n \rightarrow 5\nu)$ <span>760</span>						
See the note on $\tau(n \rightarrow 3\nu)$ on the previous data block.						

LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>0.0017	$n$	90			1 GLICENSTEIN 97	KAMI

<sup>1</sup> GLICENSTEIN 97 uses Kamioka data and the idea that the disappearance of the neutron’s magnetic moment should produce radiation.

## Inclusive modes

$\tau(N \rightarrow e^+ \text{ anything})$ <span>761</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>0.6	$p, n$	90			1 LEARNED 79	RVUE

<sup>1</sup> The electron may be primary or secondary.

$\tau(N \rightarrow \mu^+ \text{ anything})$ <span>762</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>12	$p, n$	90	2		1,2 CHERRY 81	HOME
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 1.8	$p, n$	90			2 COWSIK 80	CNTR
> 6	$p, n$	90			2 LEARNED 79	RVUE

<sup>1</sup> We have converted 2 possible events to 90% CL limit.

<sup>2</sup> The muon may be primary or secondary.

$\tau(N \rightarrow \nu \text{ anything})$ <span>763</span>						
Anything = $\pi, \rho, K$ , etc.						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>0.0002	$p, n$	90	0		LEARNED 79	RVUE

$\tau(N \rightarrow e^+ \pi^0 \text{ anything})$ <span>764</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
>0.6	$p, n$	90	0		LEARNED 79	RVUE

$\tau(N \rightarrow 2 \text{ bodies}, \nu\text{-free})$ <span>765</span>						
LIMIT ( $10^{30}$ years)	PARTICLE	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>1.3	$p, n$	90	0		ALEKSEEV 81	BAKS

## $\Delta B = 2$ dinucleon modes

$\tau(pp \rightarrow \pi^+ \pi^+)$ <span>766</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>72.2	90	2	4.45	GUSTAFSON 15	SKAM	per oxygen nucleus
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 0.7	90	4	2.34	BERGER 91B	FREJ	per iron nucleus

$\tau(pn \rightarrow \pi^+ \pi^0)$ <span>767</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>170	90			GUSTAFSON 15	SKAM	per oxygen nucleus
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 2.0	90	0	0.31	BERGER 91B	FREJ	per iron nucleus

$\tau(nn \rightarrow \pi^+ \pi^-)$ <span>768</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>0.7	90	4	2.18	BERGER 91B	FREJ	$\tau$ per iron nucleus

$\tau(nn \rightarrow \pi^0 \pi^0)$ <span>769</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>404	90			GUSTAFSON 15	SKAM	per oxygen nucleus
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 3.4	90	0	0.78	BERGER 91B	FREJ	per iron nucleus

$\tau(pp \rightarrow K^+ K^+)$ <span>770</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>170	90	0	0.28	LITOS 14	SKAM	$\tau$ per oxygen nucleus

$\tau(pp \rightarrow e^+ e^+)$ <span>771</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>5.8	90	0	<0.1	BERGER 91B	FREJ	$\tau$ per iron nucleus

$\tau(pp \rightarrow e^+ \mu^+)$ <span>772</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>3.6	90	0	<0.1	BERGER 91B	FREJ	$\tau$ per iron nucleus

$\tau(pp \rightarrow \mu^+ \mu^+)$ <span>773</span>						
LIMIT ( $10^{30}$ years)	CL%	EVTs	BKGD EST	DOCUMENT ID	TECN	COMMENT
>1.7	90	0	0.62	BERGER 91B	FREJ	$\tau$ per iron nucleus

See key on page 601

Baryon Particle Listings

*p*

$\tau(p n \rightarrow e^+ \bar{\nu})$						774
LIMIT (10 <sup>30</sup> years)	CL%	EVTS	BKGD EST	DOCUMENT ID	TECN	COMMENT
>260	90			TAKHISTOV	15	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 2.8	90	5	9.67	BERGER	91B	FREJ $\tau$ per iron nucleus

$\tau(p n \rightarrow \mu^+ \bar{\nu})$						775
LIMIT (10 <sup>30</sup> years)	CL%	EVTS	BKGD EST	DOCUMENT ID	TECN	COMMENT
>200	90			TAKHISTOV	15	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 1.6	90	4	4.37	BERGER	91B	FREJ $\tau$ per iron nucleus

$\tau(p n \rightarrow \tau^+ \bar{\nu}_\tau)$						776
LIMIT (10 <sup>30</sup> years)	CL%	EVTS	BKGD EST	DOCUMENT ID	TECN	COMMENT
>29	90			TAKHISTOV	15	SKAM
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 1	90			<sup>1</sup> BRYMAN	14	CHER
<sup>1</sup> BRYMAN 14 uses a MCGREW 99 limit on the $p \rightarrow e^+ \nu \nu$ lifetime to extract this value.						

$\tau(n n \rightarrow \nu_e \bar{\nu}_e)$						777
We include “invisible” modes here.						
LIMIT (10 <sup>30</sup> years)	CL%	EVTS	BKGD EST	DOCUMENT ID	TECN	COMMENT
>1.4	90			<sup>1</sup> ARAKI	06	KLND $nn \rightarrow$ invisible
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>0.000042	90			<sup>2</sup> TRETYAK	04	CNTR $nn \rightarrow$ invisible
>0.000049	90			<sup>3</sup> BACK	03	BORX $nn \rightarrow$ invisible
>0.000012	90			<sup>4</sup> BERNABEI	00B	DAMA $nn \rightarrow$ invisible
>0.000012	90	5	9.7	BERGER	91B	FREJ $\tau$ per iron nucleus
<sup>1</sup> ARAKI 06 looks for signs of de-excitation of the residual nucleus after disappearance of two neutrons from the s shell of <sup>12</sup> C.						
<sup>2</sup> TRETYAK 04 uses data from an old Homestake-mine radiochemical experiment on limits for invisible decays of <sup>39</sup> K to <sup>37</sup> Ar.						
<sup>3</sup> BACK 03 looks for decays of unstable nuclides left after $NN$ decays of parent <sup>12</sup> C, <sup>13</sup> C, <sup>16</sup> O nuclei. These are “invisible channel” limits.						
<sup>4</sup> BERNABEI 00b looks for the decay of a <sup>127</sup> Xe nucleus following the disappearance of an $nn$ pair in the otherwise-stable <sup>129</sup> Xe nucleus. The limit here applies as well to $nn \rightarrow \nu_\mu \bar{\nu}_\mu$ , $nn \rightarrow \nu_\tau \bar{\nu}_\tau$ , or any “disappearance” mode.						

$\tau(n n \rightarrow \nu_\mu \bar{\nu}_\mu)$							778
See the proceeding data block. "Invisible modes" would include any multi-neutrino mode.							
$LIMIT$ ( $10^{30}$ years)	CL%	EVTS	BKGD EST	CL%	DOCUMENT ID	TECN	COMMENT
>1.4	(CL = 90%) OUR LIMIT						
• • • We do not use the following data for averages, fits, limits, etc. • • •							
>0.000006	90	4	4.4		BERGER	91B	FREJ $\tau$ per iron nucleus

$(pn \rightarrow \text{invisible})$				779
This violates charge conservation as well as baryon number conservation.				
VALUE (10 <sup>30</sup> years)	CL%	DOCUMENT ID	TECN	
>0.000021	90	<sup>1</sup> TRETYAK	04	CNTR
<sup>1</sup> TRETYAK 04 uses data from an old Homestake-mine radiochemical experiment on limits for invisible decays of <sup>39</sup> K to <sup>37</sup> Ar.				

$\tau(pp \rightarrow \text{invisible})$						780
This violates charge conservation as well as baryon number conservation.						
LIMIT (10 <sup>30</sup> years)	CL%	EVTS	BKGD EST	CL%	DOCUMENT ID	TECN
>0.000005	90				<sup>1</sup> BACK	03
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>0.0000055	90				<sup>2</sup> BERNABEI	00B
<sup>1</sup> BACK 03 looks for decays of unstable nuclides left after $NN$ decays of parent <sup>12</sup> C, <sup>13</sup> C, <sup>16</sup> O nuclei. These are “invisible channel” limits.						
<sup>2</sup> BERNABEI 00b looks for the decay of a <sup>127</sup> Te nucleus following the disappearance of a $pp$ pair in the otherwise-stable <sup>129</sup> Xe nucleus.						

**$\bar{p}$  PARTIAL MEAN LIVES**

The “partial mean life” limits tabulated here are the limits on  $\bar{\tau}/B_j$ , where  $\bar{\tau}$  is the total mean life for the antiproton and  $B_j$  is the branching fraction for the mode in question.

$\tau(\bar{p} \rightarrow e^- \gamma)$						781
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
> 7 × 10 <sup>5</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>1848	95	GEER	94	CALO	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow \mu^- \gamma)$						782
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>5 × 10 <sup>4</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>5.0 × 10 <sup>4</sup>	90	HU	98B	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow e^- \pi^0)$						783
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
> 4 × 10 <sup>5</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>554	95	GEER	94	CALO	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow \mu^- \pi^0)$						784
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>5 × 10 <sup>4</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>4.8 × 10 <sup>4</sup>	90	HU	98B	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow e^- \eta)$						785
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
> 2 × 10 <sup>4</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>171	95	GEER	94	CALO	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow \mu^- \eta)$						786
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>8 × 10 <sup>3</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>7.9 × 10 <sup>3</sup>	90	HU	98B	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow e^- K_S^0)$						787
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>900	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 29	95	GEER	94	CALO	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow \mu^- K_S^0)$						788
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>4 × 10 <sup>3</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>4.3 × 10 <sup>3</sup>	90	HU	98B	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow e^- K_L^0)$						789
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>9 × 10 <sup>3</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>9	95	GEER	94	CALO	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow \mu^- K_L^0)$						790
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>7 × 10 <sup>3</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>6.5 × 10 <sup>3</sup>	90	HU	98B	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow e^- \gamma \gamma)$						791
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>2 × 10 <sup>4</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow \mu^- \gamma \gamma)$						792
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>2 × 10 <sup>4</sup>	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
>2.3 × 10 <sup>4</sup>	90	HU	98B	APEX	8.9 GeV/c $\bar{p}$ beam	

$\tau(\bar{p} \rightarrow e^- \omega)$						793
VALUE (years)	CL%	DOCUMENT ID	TECN	COMMENT		
>200	90	GEER	00	APEX	8.9 GeV/c $\bar{p}$ beam	

Baryon Particle Listings

*p*, *n*

*p* REFERENCES

MOHR 16 arXiv:1507.07956  
Accepted for publication in RMP  
ASAKURA 15 PR D92 052006  
GUSTAFSON 15 PR D91 072009  
LEE 15 PR D92 013013  
PESET 15 EPJ A51 32  
TAKHISTOV 15 PRL 115 121803  
ABE 14E PRL 113 121802  
ABE 14G PR D90 072005  
BRYMAN 14 PR B733 190  
EPSTEIN 14 PR D90 074027  
KARSHENBOIM 14A PR D90 053012  
LITOS 14 PRL 112 131803  
LORENZ 14 PL B737 57  
PDG 14 CPC 38 070001  
TAKHISTOV 14 PRL 113 101801  
ANTOGNINI 13 SCI 339 417  
DISCIACCA 13 PRL 110 130801  
MCGOVERN 13 EPJ A49 12  
MOHR 12 RMP 84 1527  
NISHINO 12 PR D85 112001  
REGIS 12 PR D86 012006  
ARRINGTON 11 PRL 107 119101  
BERNAUER 11 PRL 107 119102  
BRESSI 11 PR A83 052101  
CLOET 11 PR C83 012201  
DERUJULA 11 PL B697 26  
DISTLER 11 PL B696 343  
HILL 11 PRL 107 160402  
HORI 11 NAT 475 484  
ZHAN 11 PL B705 59  
BERNAUER 10 PRL 105 242001  
Also PR C90 015206  
BORISYUK 10 NP A843 59  
DERUJULA 10 PL B693 555  
HILL 10 PR D82 113005  
POHL 10 NAT 466 213  
NISHINO 09 PRL 102 141801  
PASK 09 PR B678 55  
MOHR 08 RMP 80 633  
BELUSHKIN 07 PR C75 038202  
ARAKI 06 PRL 96 101802  
HORI 06 PRL 96 243401  
BLUNDEN 05 PR C72 057601  
KOBAYASHI 05 PR D72 052007  
MOHR 05 RMP 77 1  
SCHUMACHER 05 PPNP 55 567  
AHMED 04 PRL 92 102004  
TRETAYAK 04 JETPL 79 106  
Translated from ZETFP 79 136  
BACK 03 PL B563 23  
BEANE 03 PL B567 200  
Also PR B607 320 (err.)  
DMITRIEV 03 PRL 91 212303  
HORI 03 PRL 91 123401  
SICK 03 PL B576 62  
ZDESENKO 03 PL B553 135  
AHMAD 02 PRL 89 011301  
BARANOV 01 PPN 32 376  
Translated from FECAY 32 699  
BLANPIED 01 PR C54 025203  
HORI 01 PRL 87 093401  
OLMOSDEL... 01 EPJ A10 207  
TRETAYAK 01 PL B505 59  
BERNABEI 00B PL B493 12  
GEER 00 PRL 84 590  
Also PR D62 052004  
Also PRL 85 3546 (err.)  
GEER 00D APJ 532 648  
SENGUPTA 00 PL B484 275  
WALL 00 PR D61 072004  
WALL 00B PR D62 092003  
GABRIELSE 00 PR 82 3198  
HAYATO 00 PR 83 1529  
MCGREW 99 PR D59 052004  
MOHR 99 JPCRD 28 1713  
Also RMP 72 351  
TORII 99 PR A59 223  
ALLISON 98 PL B427 217  
HU 98B PR D58 111101  
SHIOZAWA 98 PRL 81 3319  
GLICENSTEIN 97 PL B411 326  
GABRIELSE 95 PRL 74 3544  
MAGGIBBON 95 PR C52 2097  
GEER 94 PRL 72 1596  
HALLIN 93 PR C48 1497  
SUZUKI 93B PL B311 357  
HUGHES 92 PRL 69 578  
ZIEGER 92 PL B278 34  
Also PL B281 417 (erratum)  
BERGER 91 ZPHY C50 385  
BERGER 91B PL B269 227  
FEDERSPIEL 91 PRL 67 1511  
BECKER-SZ... 90 PR D42 2974  
ERICS ON 90 EPL 11 295  
GABRIELSE 90 PRL 65 1317  
BERGER 89 NP B313 509  
CHO 89 PRL 63 2559  
HIRATA 89C PL B220 308  
PHILLIPS 89 PL B224 348  
KREISSL 88 ZPHY C37 557  
SEIDEL 88 PRL 61 2522  
BARTELT 87 PR D36 1990  
Also PR D40 1701 (erratum)  
COHEN 87 RMP 59 1121  
HAINES 86 PRL 57 1986  
KAJITA 86 JPSJ 55 711  
ARISAKA 85 JPSJ 54 3213  
BLEWITT 85 PRL 55 2114  
DZUBA 85 PL 154B 93  
PARK 85 PRL 54 22  
BATTISTONI 84 PL 133B 454  
MARINELLI 84 PL 137B 439  
WILKENING 84 PR A29 425  
BARTELT 83 PRL 50 651  
BATTISTONI 82 PL 118B 461  
KRISHNA... 82 PL 115B 349  
ALEKSEEV 81 JETPL 33 651  
Translated from ZETFP 33 664

P.J. Mohr, D.B. Newell, B.N. Taylor (NIST)  
K. Asakura *et al.* (KamLAND Collab.)  
J. Gustafson *et al.* (Super-Kamiokande Collab.)  
G. Lee, J.R. Arrington, R.J. Hill (ANL, EFI+)  
C. Peset, A. Pineda (BARC)  
V. Takhistov *et al.* (Super-Kamiokande Collab.)  
K. Abe *et al.* (Super-Kamiokande Collab.)  
K. Abe *et al.* (Super-Kamiokande Collab.)  
D. Bryman (BRCO)  
Z. Epstein, G. Paz, J. Roy (UMD, WAYN)  
S.G. Karshenboim (MPIG)  
M. Litos *et al.* (Super-Kamiokande Collab.)  
I.T. Lorentz, U.-G. Meissner (BONN, JULI)  
K. Olive *et al.* (PDG Collab.)  
V. Takhistov *et al.* (Super-Kamiokande Collab.)  
A. Antognini *et al.* (MPIM, ETH, UPMC+)  
J. DISCIACCA *et al.* (ATRAP Collab.)  
J.A. McGovern, D.R. Phillips, H.W. Griesshammer  
P.J. Mohr, B.N. Taylor, D.B. Newell (NIST)  
H. Nishino *et al.* (Super-Kamiokande Collab.)  
C. Regis *et al.* (Super-Kamiokande Collab.)  
J. Arrington (ANL)  
J.C. Bernauer *et al.* (MAMI A1 Collab.)  
G. Bressi *et al.* (LEGN, PAVII, PADO, TRST+)  
I.C. Cloet, G.A. Miller (WASH)  
A. de Rujula (MADE, BOST, CERN)  
M.O. Distler, J.C. Bernauer, T. Walcher (MANZ)  
R.J. Hill, G. Paz (EFI)  
M. Hori *et al.* (MPIG, TOKY, BUDA, +)  
X. Zhan *et al.* (JLAB-Hall A Collab.)  
J.C. Bernauer *et al.* (MAMI A1 Collab.)  
J.C. Bernauer *et al.* (MAMI A1 Collab.)  
D. Borisjuk (KIEV)  
A. De Rujula (MADU, CERN)  
R.J. Hill, G. Paz (CHIC)  
R. Pohl *et al.* (MPIQ, ENSP, COIM, +)  
H. Nishino *et al.* (Super-Kamiokande Collab.)  
T. Pask *et al.* (Stefan Meyer Inst., Vienna, TOKY+)  
P.J. Mohr, B.N. Taylor, D.B. Newell (NIST)  
M.A. Belushkin, H.W. Hammer, U.-G. Meissner (BONN+)  
T. Araki *et al.* (JLAB-Hall A Collab.)  
H. Nishino *et al.* (CERN, TOKYO+)  
P.C. Blunden, I. Sick (MAMI, BASL)  
K. Kobayashi *et al.* (Super-Kamiokande Collab.)  
P.J. Mohr, B.N. Taylor (NIST)  
M. Schumacher (GOET)  
S.N. Ahmed *et al.* (SNO Collab.)  
V.I. Tretyak, V.Yu. Denisov, Yu.G. Zdesenko (KIEV)  
H.O. Back *et al.* (BOREXINO Collab.)  
S.R. Beane *et al.* (NOVO)  
S.R. Beane *et al.* (CERN ASACUSA Collab.)  
V.F. Dmitriev, R.A. Senkov (CERN ASACUSA Collab.)  
M. Hori *et al.* (BASIL)  
I. Sick (KIEV)  
Yu.G. Zdesenko, V.I. Tretyak (KIEV)  
Q.R. Ahmad *et al.* (SNO Collab.)  
P.S. Baranov *et al.* (SNO Collab.)

CHERRY 81 PRL 47 1507  
COWSIK 80 PR D22 2204  
BELL 79 PL 86B 215  
GOLDEN 79 PRL 43 1196  
LEARNED 79 PRL 43 907  
BREGMAN 78 PL 78B 174  
ROBERTS 78 PR D17 358  
EVANS 77 SCI 197 989  
HU 75 NP A254 403  
COHEN 73 JPCRD 2 664  
DYLLA 73 PR A7 1224  
DIX 70 Thesis Case  
HARRISON 69 PRL 22 1263  
GURR 67 PR 158 1321  
FLEROV 58 DOKL 3 79  
M.L. Cherry *et al.* (PENN, BNL)  
R. Cowsik, V.S. Narasimham (TATA)  
M. Bell *et al.* (CERN)  
R.L. Golden *et al.* (NASA, PSLL)  
J.G. Learned, F. Reines, A. Soni (UCI)  
M. Bregman *et al.* (CERN)  
B.L. Roberts (WILL, RHEL)  
J.C. Evans Jr., R.I. Steinberg (BNL, PENN)  
E. Hu *et al.* (COLU, YALE)  
E.R. Cohen, B.N. Taylor (RISC, NBS)  
H.F. Dylla, J.G. King (MIT)  
F.W. Dix (CASE)  
G.E. Harrison, P.G.H. Sanders, S.J. Wright (OXF)  
H.S. Gurr *et al.* (CASE, WITW)  
G.N. Flerov *et al.* (ASCI)

*n*  $I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \* \* \* \*

We have omitted some results that have been superseded by later experiments. See our earlier editions.

Anyone interested in the neutron should look at these two new review articles: D. Dubbers and M.G. Schmidt, "The neutron and its role in cosmology and particle physics," *Reviews of Modern Physics* **83** 1111 (2011); and F.E. Wietfeldt and G.L. Greene, "The neutron lifetime," *Reviews of Modern Physics* **83** 1173 (2011).

*n* MASS (atomic mass units u)

The mass is known much more precisely in u (atomic mass units) than in MeV. See the next data block.

VALUE (u)	DOCUMENT ID	TECN	COMMENT
<b>1.00866491588 ± 0.00000000049</b>	MOHR	16	RVUE 2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.00866491600 ± 0.00000000043	MOHR	12	RVUE 2010 CODATA value
1.00866491597 ± 0.00000000043	MOHR	08	RVUE 2006 CODATA value
1.00866491560 ± 0.00000000055	MOHR	05	RVUE 2002 CODATA value
1.00866491578 ± 0.00000000055	MOHR	99	RVUE 1998 CODATA value
1.0086655904 ± 0.0000000014	COHEN	87	RVUE 1986 CODATA value

*n* MASS (MeV)

The mass is known much more precisely in u (atomic mass units) than in MeV. The conversion from u to MeV, 1 u = 931.494 0054(57) MeV/c<sup>2</sup> (MOHR 16, the 2014 CODATA value), involves the relatively poorly known electronic charge.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>939.5654133 ± 0.0000058</b>	MOHR	16	RVUE 2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
939.565379 ± 0.000021	MOHR	12	RVUE 2010 CODATA value
939.565346 ± 0.000023	MOHR	08	RVUE 2006 CODATA value
939.565360 ± 0.000081	MOHR	05	RVUE 2002 CODATA value
939.565331 ± 0.000037	<sup>1</sup> KESSLER	99	SPEC <i>np</i> → <i>dγ</i>
939.565330 ± 0.000038	MOHR	99	RVUE 1998 CODATA value
939.56565 ± 0.00028	<sup>2,3</sup> DIFILIPPO	94	TRAP Penning trap
939.56563 ± 0.00028	COHEN	87	RVUE 1986 CODATA value
939.56564 ± 0.00028	<sup>3,4</sup> GREENE	86	SPEC <i>np</i> → <i>dγ</i>
939.5731 ± 0.0027	<sup>3</sup> COHEN	73	RVUE 1973 CODATA value

<sup>1</sup> We use the 1998 CODATA u-to-MeV conversion factor (see the heading above) to get this mass in MeV from the much more precisely measured KESSLER 99 value of 1.00866491637 ± 0.00000000082 u.  
<sup>2</sup> The mass is known much more precisely in u: *m* = 1.0086649235 ± 0.0000000023 u. We use the 1986 CODATA conversion factor to get the mass in MeV.  
<sup>3</sup> These determinations are not independent of the *m<sub>p</sub>* − *m<sub>p</sub>* measurements below.  
<sup>4</sup> The mass is known much more precisely in u: *m* = 1.008664919 ± 0.000000014 u.

*π* MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>939.485 ± 0.051</b>	59	<sup>1</sup> CRESTI	86	HBC <i>pπ</i> → <i>πn</i>

<sup>1</sup> This is a corrected result (see the erratum). The error is statistical. The maximum systematic error is 0.029 MeV.

$(m_n - m_{\pi}) / m_n$

A test of *CPT* invariance. Calculated from the *n* and *π* masses, above.

VALUE	DOCUMENT ID
<b>(9 ± 6) × 10<sup>−5</sup> OUR EVALUATION</b>	



*m<sub>n</sub> – m<sub>p</sub>*

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1.29333205 ± 0.00000051</b>	<sup>1</sup> MOHR 16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.29333217 ± 0.00000042	<sup>2</sup> MOHR 12	RVUE	2010 CODATA value
1.29333214 ± 0.00000043	<sup>3</sup> MOHR 08	RVUE	2006 CODATA value
1.2933317 ± 0.00000005	<sup>4</sup> MOHR 05	RVUE	2002 CODATA value
1.2933318 ± 0.00000005	<sup>5</sup> MOHR 99	RVUE	1998 CODATA value
1.293318 ± 0.000009	<sup>6</sup> COHEN 87	RVUE	1986 CODATA value
1.2933328 ± 0.00000072	GREENE 86	SPEC	<i>np</i> → <i>dγ</i>
1.293429 ± 0.000036	COHEN 73	RVUE	1973 CODATA value

<sup>1</sup> The 2014 CODATA mass difference in *u* is  $m_n - m_p = 1.001\,388\,449\,00(51) \times 10^{-3} u$ .

<sup>2</sup> The 2010 CODATA mass difference in *u* is  $m_n - m_p = 1.388\,449\,19(45) \times 10^{-3} u$ .

<sup>3</sup> Calculated by us from the MOHR 08 ratio  $m_n/m_p = 1.00137841918(46)$ . In *u*,  $m_n - m_p = 1.38844920(46) \times 10^{-3} u$ .

<sup>4</sup> Calculated by us from the MOHR 05 ratio  $m_n/m_p = 1.00137841870 \pm 0.00000000058$ . In *u*,  $m_n - m_p = (1.3884487 \pm 0.00000006) \times 10^{-3} u$ .

<sup>5</sup> Calculated by us from the MOHR 99 ratio  $m_n/m_p = 1.00137841887 \pm 0.00000000058$ . In *u*,  $m_n - m_p = (1.3884489 \pm 0.00000006) \times 10^{-3} u$ .

<sup>6</sup> Calculated by us from the COHEN 87 ratio  $m_n/m_p = 1.001378404 \pm 0.000000009$ . In *u*,  $m_n - m_p = 0.001388434 \pm 0.000000009 u$ .

*n* MEAN LIFE

Limits on lifetimes for *bound* neutrons are given in the section “p PARTIAL MEAN LIVES.”

We average the best seven measurements. The result,  $880.2 \pm 1.0$  s (including a scale factor of 1.9), is 5.5 seconds lower than the value we gave in 2010—a drop of 6.9 old and 5.5 new standard deviations.

For a full review of all matters concerning the neutron lifetime, see F.E. Wietfeldt and G.L. Greene, “The neutron lifetime,” *Reviews of Modern Physics* **83** 1173 (2011). In particular, there is a full discussion of the experimental methods and results; and an average lifetime is obtained making several different selections of the results then available.

VALUE (s)	DOCUMENT ID	TECN	COMMENT
<b>880.2 ± 1.0 OUR AVERAGE</b>	Error includes scale factor of 1.9. See the ideogram below.		
880.2 ± 1.2	<sup>1</sup> ARZUMANOV 15	CNTR	UCN double bottle
887.7 ± 1.2 ± 1.9	<sup>2</sup> YUE 13	CNTR	In-beam <i>n</i> , trapped <i>p</i>
882.5 ± 1.4 ± 1.5	<sup>3</sup> STEYERL 12	CNTR	UCN material bottle
880.7 ± 1.3 ± 1.2	PICHLMAIER 10	CNTR	UCN material bottle
878.5 ± 0.7 ± 0.3	SEREBROV 05	CNTR	UCN gravitational trap
889.2 ± 3.0 ± 3.8	BYRNE 96	CNTR	Penning trap
882.6 ± 2.7	<sup>4</sup> MAMPE 93	CNTR	UCN material bottle
• • • We do not use the following data for averages, fits, limits, etc. • • •			
881.6 ± 0.8 ± 1.9	<sup>5</sup> ARZUMANOV 12	CNTR	See ARZUMANOV 15
886.3 ± 1.2 ± 3.2	NICO 05	CNTR	See YUE 13
886.8 ± 1.2 ± 3.2	DEWEY 03	CNTR	See NICO 05
885.4 ± 0.9 ± 0.4	ARZUMANOV 00	CNTR	See ARZUMANOV 12
888.4 ± 3.1 ± 1.1	<sup>6</sup> NESVIZHEV... 92	CNTR	UCN material bottle
888.4 ± 2.9	ALFIMENKOV 90	CNTR	See NESVIZHEVSKII 92
893.6 ± 3.8 ± 3.7	BYRNE 90	CNTR	See BYRNE 96
878 ± 27 ± 14	KOSSAKOW... 89	TPC	Pulsed beam
887.6 ± 3.0	MAMPE 89	CNTR	See STEYERL 12
877 ± 10	PAUL 89	CNTR	Magnetic storage ring
876 ± 10 ± 19	LAST 88	SPEC	Pulsed beam
891 ± 9	SPIVAK 88	CNTR	Beam
903 ± 13	KOSVINTSEV 86	CNTR	UCN material bottle
937 ± 18	<sup>7</sup> BYRNE 80	CNTR	
875 ± 95	KOSVINTSEV 80	CNTR	
881 ± 8	BONDAREN... 78	CNTR	See SPIVAK 88
918 ± 14	CHRISTENSEN72	CNTR	

<sup>1</sup> ARZUMANOV 15 is a reanalysis of their 2008–2010 dataset, with improved systematic corrections of ARZUMANOV 00 and ARZUMANOV 12.

<sup>2</sup> YUE 13 differs from NICO 05 in that a different and better method was used to measure the neutron density in the fiducial volume. This shifted the lifetime by +1.4 seconds and reduced the previously largest source of systematic uncertainty by a factor of five.

<sup>3</sup> STEYERL 12 is a detailed reanalysis of neutron storage loss corrections to the raw data of MAMPE 89, and it replaces that value.

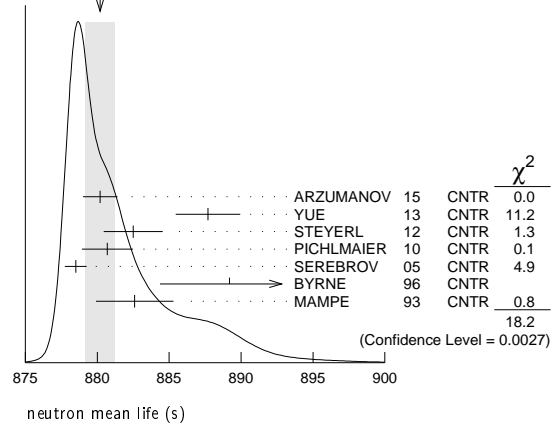
<sup>4</sup> IGNATOVICH 95 calls into question some of the corrections and averaging procedures used by MAMPE 93. The response, BONDARENKO 96, denies the validity of the criticisms.

<sup>5</sup> ARZUMANOV 12 reanalyzes its systematic corrections in ARZUMANOV 00 and obtains this corrected value.

<sup>6</sup> The NESVIZHEVSKII 92 measurement has been withdrawn by A. Serebrov.

<sup>7</sup> The BYRNE 80 measurement has been withdrawn (J. Byrne, private communication, 1990).

WEIGHTED AVERAGE  
880.2 ± 1.0 (Error scaled by 1.9)

*n* MAGNETIC MOMENT

See the “Note on Baryon Magnetic Moments” in the *A* Listings.

VALUE ( $\mu_N$ )	DOCUMENT ID	TECN	COMMENT
<b>–1.91304273 ± 0.00000045</b>	MOHR 16	RVUE	2014 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–1.91304272 ± 0.00000045	MOHR 12	RVUE	2010 CODATA value
–1.91304273 ± 0.00000045	MOHR 08	RVUE	2006 CODATA value
–1.91304273 ± 0.00000045	MOHR 05	RVUE	2002 CODATA value
–1.91304272 ± 0.00000045	MOHR 99	RVUE	1998 CODATA value
–1.91304275 ± 0.00000045	COHEN 87	RVUE	1986 CODATA value
–1.91304277 ± 0.00000048	<sup>1</sup> GREENE 82	MRS	

<sup>1</sup> GREENE 82 measures the moment to be  $(1.04187564 \pm 0.00000026) \times 10^{-3}$  Bohr magnetons. The value above is obtained by multiplying this by  $m_p/m_e = 1836.152701 \pm 0.000037$  (the 1986 CODATA value from COHEN 87).

*n* ELECTRIC DIPOLE MOMENT

A nonzero value is forbidden by both *T* invariance and *P* invariance. A number of early results have been omitted. See RAMSEY 90, GOLUB 94, and LAMOREAUX 09 for reviews.

The results are upper limits on  $|d_n|$ .

VALUE ( $10^{-25}$ ecm)	CL %	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.30</b>	90	PENDLEBURY 15	MRS	$d = (-0.21 \pm 1.82) \times 10^{-26}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.55	90	SEREBROV 15	MRS	UCN's, $h\nu = 2\mu_n B \pm 2d_n E$
< 0.55	90	<sup>1</sup> SEREBROV 14	MRS	See SEREBROV 15
< 0.29	90	<sup>2</sup> BAKER 06	MRS	See PENDLEBURY 15
< 0.63	90	<sup>3</sup> HARRIS 99	MRS	$d = (-0.1 \pm 0.36) \times 10^{-25}$
< 0.97	90	ALTAREV 96	MRS	See SEREBROV 14
< 1.1	95	ALTAREV 92	MRS	See ALTAREV 96
< 1.2	95	SMITH 90	MRS	See HARRIS 99
< 2.6	95	ALTAREV 86	MRS	$d = (-1.4 \pm 0.6) \times 10^{-25}$
0.3 ± 4.8		PENDLEBURY 84	MRS	Ultracold neutrons
< 6	90	ALTAREV 81	MRS	$d = (2.1 \pm 2.4) \times 10^{-25}$
< 16	90	ALTAREV 79	MRS	$d = (4.0 \pm 7.5) \times 10^{-25}$

<sup>1</sup> SEREBROV 14 includes the data of ALTAREV 96.

<sup>2</sup> LAMOREAUX 07 faults BAKER 06 for not including in the estimate of systematic error an effect due to the Earth's rotation. BAKER 07 replies (1) that the effect was included implicitly in the analysis and (2) that further analysis confirms that the BAKER 06 limit is correct as is. See also SILENKO 07.

<sup>3</sup> This HARRIS 99 result includes the result of SMITH 90. However, the averaging of the results of these two experiments has been criticized by LAMOREAUX 00.

*n* MEAN-SQUARE CHARGE RADIUS

The mean-square charge radius of the neutron,  $\langle r_n^2 \rangle$ , is related to the neutron-electron scattering length  $b_{ne}$  by  $\langle r_n^2 \rangle = 3(m_e a_0 / m_n) b_{ne}$ , where  $m_e$  and  $m_n$  are the masses of the electron and neutron, and  $a_0$  is the Bohr radius. Numerically,  $\langle r_n^2 \rangle = 86.34 b_{ne}$ , if we use  $a_0$  for a nucleus with infinite mass.

VALUE ( $\text{fm}^2$ )	DOCUMENT ID	COMMENT
<b>–0.1161 ± 0.0022 OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.	
–0.115 ± 0.002 ± 0.003	KOPECKY 97	<i>ne</i> scattering (Pb)
–0.124 ± 0.003 ± 0.005	KOPECKY 97	<i>ne</i> scattering (Bi)
–0.114 ± 0.003	KOESTER 95	<i>ne</i> scattering (Pb, Bi)
–0.134 ± 0.009	ALEKSANDR... 86	<i>ne</i> scattering (Bi)
–0.115 ± 0.003	<sup>1</sup> KROHN 73	<i>ne</i> scattering (Ne, Ar, Kr, Xe)

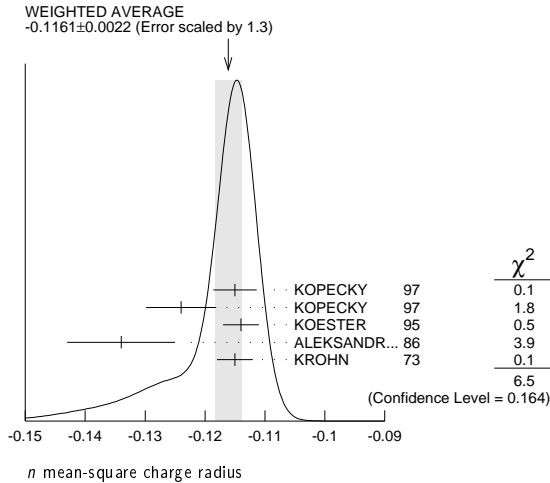
## Baryon Particle Listings

 $n$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.117 \pm 0.007$ $-0.011$	BELUSHKIN	07	Dispersion analysis
$-0.113 \pm 0.003 \pm 0.004$	KOPECKY	95	$ne$ scattering (Pb)
$-0.114 \pm 0.003$	KOESTER	86	$ne$ scattering (Pb, Bi)
$-0.118 \pm 0.002$	KOESTER	76	$ne$ scattering (Pb)
$-0.120 \pm 0.002$	KOESTER	76	$ne$ scattering (Bi)
$-0.116 \pm 0.003$	KROHN	66	$ne$ scattering (Ne, Ar, Kr, Xe)

<sup>1</sup> This value is as corrected by KOESTER 76.

 $n$  MAGNETIC RADIUS

This is the rms magnetic radius,  $\sqrt{\langle r_M^2 \rangle}$ .

VALUE (fm)	DOCUMENT ID	COMMENT
<b>0.864 ± 0.009</b> <b>-0.008</b> OUR AVERAGE		
0.89 ± 0.03	EPSTEIN 14	Using $ep$ , $en$ , $\pi\pi$ data
0.862 ± 0.009 -0.008	BELUSHKIN 07	Dispersion analysis

 $n$  ELECTRIC POLARIZABILITY  $\alpha_n$ 

Following is the electric polarizability  $\alpha_n$  defined in terms of the induced electric dipole moment by  $\mathbf{D} = 4\pi\epsilon_0\alpha_n\mathbf{E}$ . For a review, see SCHMIED-MAYER 89.

For very complete reviews of the polarizability of the nucleon and Compton scattering, see SCHUMACHER 05 and GRIESSHAMMER 12.

VALUE ( $10^{-4}$ fm <sup>3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>11.8 ± 1.1</b> OUR AVERAGE			
11.55 ± 1.25 ± 0.8	MYERS 14	CNTR	$\gamma d \rightarrow \gamma d$
12.5 ± 1.8 +1.6 -1.3	<sup>1</sup> KOSSERT 03	CNTR	$\gamma d \rightarrow \gamma pn$
12.0 ± 1.5 ± 2.0	SCHMIEDM... 91	CNTR	$n$ Pb transmission
10.7 ± 3.3 -10.7	ROSE 90B	CNTR	$\gamma d \rightarrow \gamma np$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
8.8 ± 2.4 ± 3.0	<sup>2</sup> LUNDIN 03	CNTR	$\gamma d \rightarrow \gamma d$
13.6	<sup>3</sup> KOLB 00	CNTR	$\gamma d \rightarrow \gamma np$
0.0 ± 5.0	<sup>4</sup> KOESTER 95	CNTR	$n$ Pb, $n$ Bi transmission
11.7 ± 4.3 -11.7	ROSE 90	CNTR	See ROSE 90B
8 ± 10	KOESTER 88	CNTR	$n$ Pb, $n$ Bi transmission
12 ± 10	SCHMIEDM... 88	CNTR	$n$ Pb, $n$ C transmission

<sup>1</sup> KOSSERT 03 gets  $\alpha_n - \beta_n = (9.8 \pm 3.6 \pm 2.1 \pm 2.2) \times 10^{-4}$  fm<sup>3</sup>, and uses  $\alpha_n + \beta_n = (15.2 \pm 0.5) \times 10^{-4}$  fm<sup>3</sup> from LEVCHUK 00. Thus the errors on  $\alpha_n$  and  $\beta_n$  are anti-correlated.

<sup>2</sup> LUNDIN 03 measures  $\alpha_N - \beta_N = (6.4 \pm 2.4) \times 10^{-4}$  fm<sup>3</sup> and uses accurate values for  $\alpha_p$  and  $\alpha_p$  and a precise sum-rule result for  $\alpha_n + \beta_n$ . The second error is a model uncertainty, and errors on  $\alpha_n$  and  $\beta_n$  are anticorrelated. The data from this paper are included in the analysis of MYERS 14.

<sup>3</sup> KOLB 00 obtains this value with a lower limit of  $7.6 \times 10^{-4}$  fm<sup>3</sup> but no upper limit from this experiment alone. Combined with results of ROSE 90, the 1- $\sigma$  range is (7.6–14.0)  $\times 10^{-4}$  fm<sup>3</sup>.

<sup>4</sup> KOESTER 95 uses natural Pb and the isotopes 208, 207, and 206. See this paper for a discussion of methods used by various groups to extract  $\alpha_n$  from data.

 $n$  MAGNETIC POLARIZABILITY  $\beta_n$ 

VALUE ( $10^{-4}$ fm <sup>3</sup> )	DOCUMENT ID	TECN	COMMENT
<b>3.7 ± 1.2</b> OUR AVERAGE			
3.65 ± 1.25 ± 0.8	MYERS 14	CNTR	$\gamma d \rightarrow \gamma d$
2.7 ± 1.8 +1.3 -1.6	<sup>1</sup> KOSSERT 03	CNTR	$\gamma d \rightarrow \gamma pn$
6.5 ± 2.4 ± 3.0	<sup>2</sup> LUNDIN 03	CNTR	$\gamma d \rightarrow \gamma d$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.6	<sup>3</sup> KOLB 00	CNTR	$\gamma d \rightarrow \gamma np$
<sup>1</sup> KOSSERT 03 gets $\alpha_n - \beta_n = (9.8 \pm 3.6 \pm 2.1 \pm 2.2) \times 10^{-4}$ fm <sup>3</sup> , and uses $\alpha_n + \beta_n = (15.2 \pm 0.5) \times 10^{-4}$ fm <sup>3</sup> from LEVCHUK 00. Thus the errors on $\alpha_n$ and $\beta_n$ are anti-correlated.			
<sup>2</sup> LUNDIN 03 measures $\alpha_N - \beta_N = (6.4 \pm 2.4) \times 10^{-4}$ fm <sup>3</sup> and uses accurate values for $\alpha_p$ and $\alpha_p$ and a precise sum-rule result for $\alpha_n + \beta_n$ . The second error is a model uncertainty, and errors on $\alpha_n$ and $\beta_n$ are anticorrelated.			
<sup>3</sup> KOLB 00 obtains this value with an upper limit of $7.6 \times 10^{-4}$ fm <sup>3</sup> but no lower limit from this experiment alone. Combined with results of ROSE 90, the 1- $\sigma$ range is (1.2–7.6) $\times 10^{-4}$ fm <sup>3</sup> .			

 $n$  CHARGE

See also “ $|q_p + q_e|/e$ ” in the proton Listings.

VALUE ( $10^{-21}$ e)	DOCUMENT ID	TECN	COMMENT
<b>-0.2 ± 0.8</b> OUR AVERAGE			
-0.1 ± 1.1	<sup>1</sup> BRESSI 11		Neutrality of SF <sub>6</sub>
-0.4 ± 1.1	<sup>2</sup> BAUMANN 88		Cold $n$ deflection
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-15 ± 22	<sup>3</sup> GAEHLER 82	CNTR	Cold $n$ deflection
<sup>1</sup> As a limit, this BRESSI 11 value is $< 1 \times 10^{-21}$ e.			
<sup>2</sup> The BAUMANN 88 error $\pm 1.1$ gives the 68% CL limits about the the value -0.4.			
<sup>3</sup> The GAEHLER 82 error $\pm 22$ gives the 90% CL limits about the the value -15.			

LIMIT ON  $n\bar{n}$  OSCILLATIONSMean Time for  $n\bar{n}$  Transition in Vacuum

A test of  $\Delta B=2$  baryon number nonconservation. MOHAPATRA 80 and MOHAPATRA 89 discuss the theoretical motivations for looking for  $n\bar{n}$  oscillations. DOVER 83 and DOVER 85 give phenomenological analyses. The best limits come from looking for the decay of neutrons bound in nuclei. However, these analyses require model-dependent corrections for nuclear effects. See KABIR 83, DOVER 89, ALBERICO 91, and GAL 00 for discussions. Direct searches for  $n \rightarrow \bar{n}$  transitions using reactor neutrons are cleaner but give somewhat poorer limits. We include limits for both free and bound neutrons in the Summary Table. See MOHAPATRA 09 for a recent review.

VALUE (s)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;2.7 × 10<sup>8</sup></b>	90	ABE 15c	CNTR	$n$ bound in oxygen
<b>&gt;8.6 × 10<sup>7</sup></b>	90	BALDO-...	94	CNTR Reactor (free) neutrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>1.3 × 10 <sup>8</sup>	90	CHUNG 02b	SOU2	$n$ bound in iron
>1 × 10 <sup>7</sup>	90	BALDO-...	90	CNTR See BALDO-CEOLIN 94
>1.2 × 10 <sup>8</sup>	90	BERGER 90	FREJ	$n$ bound in iron
>4.9 × 10 <sup>5</sup>	90	BRESSI 90	CNTR	Reactor neutrons
>4.7 × 10 <sup>5</sup>	90	BRESSI 89	CNTR	See BRESSI 90
>1.2 × 10 <sup>8</sup>	90	TAKITA 86	CNTR	$n$ bound in oxygen
>1 × 10 <sup>6</sup>	90	FIDECARO 85	CNTR	Reactor neutrons
>8.8 × 10 <sup>7</sup>	90	PARK 85b	CNTR	
>3 × 10 <sup>7</sup>		BATTISTONI 84	NUSX	
>0.27–1.1 × 10 <sup>8</sup>		JONES 84	CNTR	
>2 × 10 <sup>7</sup>		CHERRY 83	CNTR	

LIMIT ON  $n n'$  OSCILLATIONS

Lee and Yang (LEE 56) proposed the existence of mirror world in an attempt to restore global parity symmetry. See BEREZHIANI 06 for a recent discussion.

VALUE (s)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;414</b>	90	SEREBROV 08	CNTR	UCN, B field on & off
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 12	95	<sup>1</sup> ALTAREV 09a	CNTR	UCN, scan 0 ≤ B ≤ 12.5 μT
>103	95	BAN 07	CNTR	UCN, B field on & off

<sup>1</sup> Losses of neutrons due to oscillations to mirror neutrons would be maximal when the magnetic fields  $B$  and  $B'$  in the two worlds were equal. Hence the scan over  $B$  by ALTAREV 09a: the limit applies for any  $B'$  over the given range. At  $B' = 0$ , the limit is 141 s (95% CL).

## n DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $p e^- \bar{\nu}_e$	100	%
$\Gamma_2$ $p e^- \bar{\nu}_e \gamma$	[a] $(3.09 \pm 0.32) \times 10^{-3}$	
$\Gamma_3$ hydrogen-atom $\bar{\nu}_e$		
<b>Charge conservation (Q) violating mode</b>		
$\Gamma_4$ $p \nu_e \bar{\nu}_e$	Q < 8	$\times 10^{-27}$ 68%

[a] This limit is for  $\gamma$  energies between 15 and 340 keV.

## n BRANCHING RATIOS

$\Gamma(p e^- \bar{\nu}_e \gamma)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE (units $10^{-3}$ )	CL%
<b><math>3.09 \pm 0.11 \pm 0.30</math></b>	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$3.13 \pm 0.11 \pm 0.33$	90
<6.9	90

<sup>1</sup> This COOPER 10 result is for  $\gamma$  energies between 15 and 340 keV.  
<sup>2</sup> This BECK 02 limit is for  $\gamma$  energies between 35 and 100 keV.

$\Gamma(\text{hydrogen-atom } \bar{\nu}_e)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE	CL%
• • • We do not use the following data for averages, fits, limits, etc. • • •	
< $3 \times 10^{-2}$	95

<sup>1</sup> GREEN 90 infers that  $\tau(\text{hydrogen-atom } \bar{\nu}_e) > 3 \times 10^4$  s by comparing neutron lifetime measurements made in storage experiments with those made in  $\beta$ -decay experiments. However, the result depends sensitively on the lifetime measurements, and does not of course take into account more recent measurements of same.

$\Gamma(p \nu_e \bar{\nu}_e)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
VALUE	CL%
<b>&lt;8 <math>\times 10^{-27}</math></b>	68
• • • We do not use the following data for averages, fits, limits, etc. • • •	
< $9.7 \times 10^{-18}$	90
< $7.9 \times 10^{-21}$	90
< $9 \times 10^{-24}$	90
< $3 \times 10^{-19}$	90

Forbidden by charge conservation.

<sup>1</sup> NORMAN 96 gets this limit by attributing SAGE and GALLEX counting rates to the charge-nonconserving transition  $^{71}\text{Ga} \rightarrow ^{71}\text{Ge} + \text{neutrals}$  rather than to solar-neutrino reactions.

## BARYON DECAY PARAMETERS

Written 1996 by E.D. Commins (University of California, Berkeley).

## Baryon semileptonic decays

The typical spin-1/2 baryon semileptonic decay is described by a matrix element, the hadronic part of which may be written as:

$$\bar{B}_f [f_1(q^2)\gamma_\lambda + i f_2(q^2)\sigma_{\lambda\mu}q^\mu + g_1(q^2)\gamma_\lambda\gamma_5 + g_3(q^2)\gamma_5q_\lambda] B_i. \quad (1)$$

Here  $B_i$  and  $\bar{B}_f$  are spinors describing the initial and final baryons, and  $q = p_i - p_f$ , while the terms in  $f_1$ ,  $f_2$ ,  $g_1$ , and  $g_3$  account for vector, induced tensor (“weak magnetism”), axial vector, and induced pseudoscalar contributions [1]. Second-class current contributions are ignored here. In the limit of zero momentum transfer,  $f_1$  reduces to the vector coupling constant  $g_V$ , and  $g_1$  reduces to the axial-vector coupling constant  $g_A$ . The latter coefficients are related by Cabibbo’s theory [2], generalized to six quarks (and three mixing angles) by Kobayashi and Maskawa [3]. The  $g_3$  term is negligible for transitions in which an  $e^\pm$  is emitted, and gives a very small correction, which can be estimated by PCAC [4], for  $\mu^\pm$  modes. Recoil effects

include weak magnetism, and are taken into account adequately by considering terms of first order in

$$\delta = \frac{m_i - m_f}{m_i + m_f}, \quad (2)$$

where  $m_i$  and  $m_f$  are the masses of the initial and final baryons.

The experimental quantities of interest are the total decay rate, the lepton-neutrino angular correlation, the asymmetry coefficients in the decay of a polarized initial baryon, and the polarization of the decay baryon in its own rest frame for an unpolarized initial baryon. Formulae for these quantities are derived by standard means [5] and are analogous to formulae for nuclear beta decay [6]. We use the notation of Ref. 6 in the Listings for neutron beta decay. For comparison with experiments at higher  $q^2$ , it is necessary to modify the form factors at  $q^2 = 0$  by a “dipole”  $q^2$  dependence, and for high-precision comparisons to apply appropriate radiative corrections [7].

The ratio  $g_A/g_V$  may be written as

$$g_A/g_V = |g_A/g_V| e^{i\phi_{AV}}. \quad (3)$$

The presence of a “triple correlation” term in the transition probability, proportional to  $\text{Im}(g_A/g_V)$  and of the form

$$\sigma_i \cdot (\mathbf{p}_\ell \times \mathbf{p}_\nu) \quad (4)$$

for initial baryon polarization or

$$\sigma_f \cdot (\mathbf{p}_\ell \times \mathbf{p}_\nu) \quad (5)$$

for final baryon polarization, would indicate failure of time-reversal invariance. The phase angle  $\phi$  has been measured precisely only in neutron decay (and in  $^{19}\text{Ne}$  nuclear beta decay), and the results are consistent with  $T$  invariance.

## Hyperon nonleptonic decays

The amplitude for a spin-1/2 hyperon decaying into a spin-1/2 baryon and a spin-0 meson may be written in the form

$$M = G_F m_\pi^2 \cdot \bar{B}_f (A - B\gamma_5) B_i, \quad (6)$$

where  $A$  and  $B$  are constants [1]. The transition rate is proportional to

$$R = 1 + \gamma \hat{\omega}_f \cdot \hat{\omega}_i + (1 - \gamma)(\hat{\omega}_f \cdot \hat{\mathbf{n}})(\hat{\omega}_i \cdot \hat{\mathbf{n}}) + \alpha(\hat{\omega}_f \cdot \hat{\mathbf{n}} + \hat{\omega}_i \cdot \hat{\mathbf{n}}) + \beta \hat{\mathbf{n}} \cdot (\hat{\omega}_f \times \hat{\omega}_i), \quad (7)$$

where  $\hat{\mathbf{n}}$  is a unit vector in the direction of the final baryon momentum, and  $\hat{\omega}_i$  and  $\hat{\omega}_f$  are unit vectors in the directions of the initial and final baryon spins. (The sign of the last term in the above equation was incorrect in our 1988 and 1990 editions.) The parameters  $\alpha$ ,  $\beta$ , and  $\gamma$  are defined as

$$\begin{aligned} \alpha &= 2 \text{Re}(s^*p)/(|s|^2 + |p|^2), \\ \beta &= 2 \text{Im}(s^*p)/(|s|^2 + |p|^2), \\ \gamma &= (|s|^2 - |p|^2)/(|s|^2 + |p|^2), \end{aligned} \quad (8)$$

# Baryon Particle Listings

$n$

where  $s = A$  and  $p = |\mathbf{p}_f| B/(E_f + m_f)$ ; here  $E_f$  and  $\mathbf{p}_f$  are the energy and momentum of the final baryon. The parameters  $\alpha$ ,  $\beta$ , and  $\gamma$  satisfy

$$\alpha^2 + \beta^2 + \gamma^2 = 1. \quad (9)$$

If the hyperon polarization is  $\mathbf{P}_Y$ , the polarization  $\mathbf{P}_B$  of the decay baryons is

$$\mathbf{P}_B = \frac{(\alpha + \mathbf{P}_Y \cdot \hat{\mathbf{n}})\hat{\mathbf{n}} + \beta(\mathbf{P}_Y \times \hat{\mathbf{n}}) + \gamma\hat{\mathbf{n}} \times (\mathbf{P}_Y \times \hat{\mathbf{n}})}{1 + \alpha\mathbf{P}_Y \cdot \hat{\mathbf{n}}}. \quad (10)$$

Here  $\mathbf{P}_B$  is defined in the rest system of the baryon, obtained by a Lorentz transformation along  $\hat{\mathbf{n}}$  from the hyperon rest frame, in which  $\hat{\mathbf{n}}$  and  $\mathbf{P}_Y$  are defined.

An additional useful parameter  $\phi$  is defined by

$$\beta = (1 - \alpha^2)^{1/2} \sin\phi. \quad (11)$$

In the Listings, we compile  $\alpha$  and  $\phi$  for each decay, since these quantities are most closely related to experiment and are essentially uncorrelated. When necessary, we have changed the signs of reported values to agree with our sign conventions. In the Baryon Summary Table, we give  $\alpha$ ,  $\phi$ , and  $\Delta$  (defined below) with errors, and also give the value of  $\gamma$  without error.

Time-reversal invariance requires, in the absence of final-state interactions, that  $s$  and  $p$  be relatively real, and therefore that  $\beta = 0$ . However, for the decays discussed here, the final-state interaction is strong. Thus

$$s = |s|e^{i\delta_s} \text{ and } p = |p|e^{i\delta_p}, \quad (12)$$

where  $\delta_s$  and  $\delta_p$  are the pion-baryon  $s$ - and  $p$ -wave strong interaction phase shifts. We then have

$$\beta = \frac{-2|s||p|}{|s|^2 + |p|^2} \sin(\delta_s - \delta_p). \quad (13)$$

One also defines  $\Delta = -\tan^{-1}(\beta/\alpha)$ . If  $T$  invariance holds,  $\Delta = \delta_s - \delta_p$ . For  $\Lambda \rightarrow p\pi^-$  decay, the value of  $\Delta$  may be compared with the  $s$ - and  $p$ -wave phase shifts in low-energy  $\pi^-p$  scattering, and the results are consistent with  $T$  invariance.

See also the note on "Radiative Hyperon Decays" in the  $\Xi^0$  Listings in this *Review*.

## References

1. E.D. Commins and P.H. Bucksbaum, *Weak Interactions of Leptons and Quarks* (Cambridge University Press, Cambridge, England, 1983).
2. N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
3. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
4. M.L. Goldberger and S.B. Treiman, Phys. Rev. **111**, 354 (1958).
5. P.H. Frampton and W.K. Tung, Phys. Rev. **D3**, 1114 (1971).
6. J.D. Jackson, S.B. Treiman, and H.W. Wyld, Jr., Phys. Rev. **106**, 517 (1957), and Nucl. Phys. **4**, 206 (1957).
7. Y. Yokoo, S. Suzuki, and M. Morita, Prog. Theor. Phys. **50**, 1894 (1973).

## $n \rightarrow p e^- \bar{\nu}_e$ DECAY PARAMETERS

See the above "Note on Baryon Decay Parameters." For discussions of recent results, see the references cited at the beginning of the section on the neutron mean life. For discussions of the values of the weak coupling constants  $g_A$  and  $g_V$  obtained using the neutron lifetime and asymmetry parameter  $A$ , comparisons with other methods of obtaining these constants, and implications for particle physics and for astrophysics, see DUBBERS 91 and WOOLCOCK 91. For tests of the  $V-A$  theory of neutron decay, see EROZOLIMSKII 91B, MOSTOVOI 96, NICO 05, SEVERIJNS 06, and ABELE 08.

### $\lambda \equiv g_A / g_V$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-1.2723 ± 0.0023 OUR AVERAGE</b>	Error includes scale factor of 2.2. See the ideogram below.		
-1.2755 ± 0.0030	1 MENDENHALL 13	UCNA	Ultracold $n$ , polarized
-1.2748 ± 0.0008 + 0.0010 -0.0011	2 MUND 13	SPEC	Cold $n$ , polarized
-1.275 ± 0.006 ± 0.015	SCHUMANN 08	CNTR	Cold $n$ , polarized
-1.2686 ± 0.0046 ± 0.0007	3 MOSTOVOI 01	CNTR	$A$ and $B \times$ polarizations
-1.266 ± 0.004	LIAUD 97	TPC	Cold $n$ , polarized, $A$
-1.2594 ± 0.0038	4 YEROZLIM... 97	CNTR	Cold $n$ , polarized, $A$
-1.262 ± 0.005	BOPP 86	SPEC	Cold $n$ , polarized, $A$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-1.27590 ± 0.00239 ± 0.00331 -0.00377	5 PLASTER 12	UCNA	See MENDENHALL 13
-1.27590 ± 0.00409 -0.00445	LIU 10	UCNA	See PLASTER 12
-1.2739 ± 0.0019	6 ABELE 02	SPEC	See MUND 13
-1.274 ± 0.003	ABELE 97D	SPEC	Cold $n$ , polarized, $A$
-1.266 ± 0.004	SCHRECK... 95	TPC	See LIAUD 97
-1.2544 ± 0.0036	EROZOLIM... 91	CNTR	See YEROZOLIMSKY 97
-1.226 ± 0.042	MOSTOVOY 83	RVUE	
-1.261 ± 0.012	EROZOLIM... 79	CNTR	Cold $n$ , polarized, $A$
-1.259 ± 0.017	7 STRATOWA 78	CNTR	$p$ recoil spectrum, $A$
-1.263 ± 0.015	EROZOLIM... 77	CNTR	See EROZOLIMSKII 79
-1.250 ± 0.036	7 DOBROZE... 75	CNTR	See STRATOWA 78
-1.258 ± 0.015	8 KROHN 75	CNTR	Cold $n$ , polarized, $A$
-1.263 ± 0.016	9 KROPF 74	RVUE	$n$ decay alone
-1.250 ± 0.009	9 KROPF 74	RVUE	$n$ decay + nuclear ft

1 MENDENHALL 13 gets  $A = -0.11954 \pm 0.00055 \pm 0.00098$  and  $\lambda = -1.2756 \pm 0.0030$ . We quote the nearly identical values that include the earlier UCNA measurement (PLASTER 12), with a correction to that result.

2 This MUND 13 value includes earlier PERKEO II measurements (ABELE 02 and ABELE 97D).

3 MOSTOVOI 01 measures the two  $P$ -odd correlations  $A$  and  $B$ , or rather  $SA$  and  $SB$ , where  $S$  is the  $n$  polarization, in free neutron decay.

4 YEROZOLIMSKY 97 makes a correction to the EROZOLIMSKII 91 value.

5 This PLASTER 12 value is identical with that given in LIU 10, but the experiment is now described in detail.

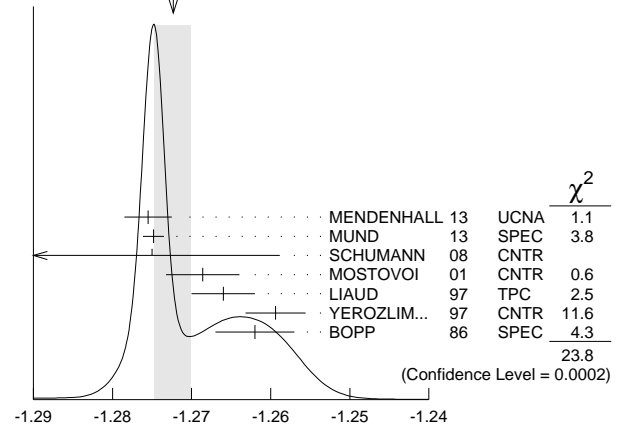
6 This is the combined result of ABELE 02 and ABELE 97D.

7 These experiments measure the absolute value of  $g_A/g_V$  only.

8 KROHN 75 includes events of CHRISTENSEN 70.

9 KROPF 74 reviews all data through 1972.

WEIGHTED AVERAGE  
-1.2723±0.0023 (Error scaled by 2.2)



$\lambda \equiv g_A / g_V$

See key on page 601

# Baryon Particle Listings

*n*

## $e^-$ ASYMMETRY PARAMETER $A$

This is the neutron-spin electron-momentum correlation coefficient. Unless otherwise noted, the values are corrected for radiative effects and weak magnetism. In the Standard Model,  $A$  is related to  $\lambda \equiv g_A/g_V$  by  $A = -2\lambda(\lambda+1)/(1+3\lambda^2)$ ; this assumes that  $g_A$  and  $g_V$  are real.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.1184 ± 0.0010 OUR AVERAGE</b>			Error includes scale factor of 2.4. See the ideogram below.
-0.11952 ± 0.00110	<sup>1</sup> MENDENHALL 13	UCNA	Ultracold <i>n</i> , polarized
-0.11926 ± 0.00031 ± 0.00036 -0.00042	<sup>2</sup> MUND 13	SPEC	Cold <i>n</i> , polarized
-0.1160 ± 0.0009 ± 0.0012	LIAUD 97	TPC	Cold <i>n</i> , polarized
-0.1135 ± 0.0014	<sup>3</sup> YEROZLIM... 97	CNTR	Cold <i>n</i> , polarized
-0.1146 ± 0.0019	BOPP 86	SPEC	Cold <i>n</i> , polarized
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.11966 ± 0.00089 ± 0.00123 -0.00140	<sup>4</sup> PLASTER 12	UCNA	See MENDENHALL 13
-0.11966 ± 0.00089 ± 0.00123 -0.00140	LIU 10	UCNA	See PLASTER 12
-0.1138 ± 0.0046 ± 0.0021	PATTIE 09	SPEC	Ultracold <i>n</i> , polarized
-0.1189 ± 0.0007	<sup>5</sup> ABELE 02	SPEC	See MUND 13
-0.1168 ± 0.0017	<sup>6</sup> MOSTOVOI 01	CNTR	Inferred
-0.1189 ± 0.0012	ABELE 97D	SPEC	Cold <i>n</i> , polarized
-0.1160 ± 0.0009 ± 0.0011	SCHRECK... 95	TPC	See LIAUD 97
-0.1116 ± 0.0014	EROZOLIM... 91	CNTR	See YEROZOLIM-SKY 97
-0.114 ± 0.005	<sup>7</sup> EROZOLIM... 79	CNTR	Cold <i>n</i> , polarized
-0.113 ± 0.006	<sup>7</sup> KROHN 75	CNTR	Cold <i>n</i> , polarized

<sup>1</sup> MENDENHALL 13 gets  $A = -0.11954 \pm 0.00055 \pm 0.00098$  and  $\lambda = -1.2756 \pm 0.0030$ . We quote the nearly identical values that include the earlier UCNA measurement (PLASTER 12), with a correction to that result.

<sup>2</sup> This MUND 13 value includes earlier PERKEO II measurements (ABELE 02 and ABELE 97D), with a correction to those results.

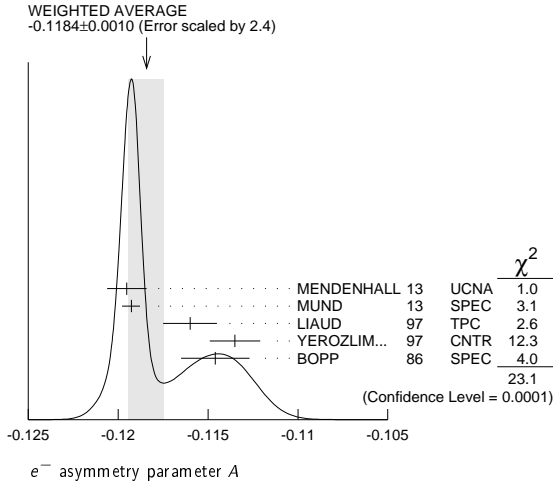
<sup>3</sup> YEROZOLIMSKY 97 makes a correction to the EROZOLIMSKII 91 value.

<sup>4</sup> This PLASTER 12 value is identical with that given in LIU 10, but the experiment is now described in detail.

<sup>5</sup> This is the combined result of ABELE 02 and ABELE 97D.

<sup>6</sup> MOSTOVOI 01 calculates this from its measurement of  $\lambda = g_A/g_V$  above.

<sup>7</sup> These results are not corrected for radiative effects and weak magnetism, but the corrections are small compared to the errors.



## $\bar{\nu}_e$ ASYMMETRY PARAMETER $B$

This is the neutron-spin antineutrino-momentum correlation coefficient. In the Standard Model,  $B$  is related to  $\lambda \equiv g_A/g_V$  by  $B = 2\lambda(\lambda-1)/(1+3\lambda^2)$ ; this assumes that  $g_A$  and  $g_V$  are real.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.9807 ± 0.0030 OUR AVERAGE</b>			
0.9802 ± 0.0034 ± 0.0036	SCHUMANN 07	CNTR	Cold <i>n</i> , polarized
0.967 ± 0.006 ± 0.010	KREUZ 05	CNTR	Cold <i>n</i> , polarized
0.9801 ± 0.0046	SEREBROV 98	CNTR	Cold <i>n</i> , polarized
0.9894 ± 0.0083	KUZNETSOV 95	CNTR	Cold <i>n</i> , polarized
1.00 ± 0.05	CHRISTENSEN70	CNTR	Cold <i>n</i> , polarized
0.995 ± 0.034	EROZOLIM... 70C	CNTR	Cold <i>n</i> , polarized
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.9876 ± 0.0004	<sup>1</sup> MOSTOVOI 01	CNTR	Inferred

<sup>1</sup> MOSTOVOI 01 calculates this from its measurement of  $\lambda = g_A/g_V$  above.

## PROTON ASYMMETRY PARAMETER $C$

Describes the correlation between the neutron spin and the proton momentum. In the Standard Model,  $C$  is related to  $\lambda \equiv g_A/g_V$  by  $C = -x_c(A+B) = x_c 4\lambda/(1+3\lambda^2)$ , where  $x_c = 0.27484$  is a kinematic factor; this assumes that  $g_A$  and  $g_V$  are real.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.2377 ± 0.0010 ± 0.0024</b>	SCHUMANN 08	CNTR	Cold <i>n</i> , polarized

## $e^- \bar{\nu}_e$ ANGULAR CORRELATION COEFFICIENT $a$

For a review of past experiments and plans for future measurements of the  $a$  parameter, see WIETFELDT 05. In the Standard Model,  $a$  is related to  $\lambda \equiv g_A/g_V$  by  $a = (1 - \lambda^2)/(1 + 3\lambda^2)$ ; this assumes that  $g_A$  and  $g_V$  are real.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.103 ± 0.004 OUR AVERAGE</b>			
-0.1054 ± 0.0055	BYRNE 02	SPEC	Proton recoil spectrum
-0.1017 ± 0.0051	STRATOWA 78	CNTR	Proton recoil spectrum
-0.091 ± 0.039	GRIGOREV 68	SPEC	Proton recoil spectrum
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.1045 ± 0.0014	<sup>1</sup> MOSTOVOI 01	CNTR	Inferred

<sup>1</sup> MOSTOVOI 01 calculates this from its measurement of  $\lambda = g_A/g_V$  above.

## $\phi_{AV}$ PHASE OF $g_A$ RELATIVE TO $g_V$

Time reversal invariance requires this to be 0 or 180°. This is related to  $D$  given in the next data block and  $\lambda \equiv g_A/g_V$  by  $\sin(\phi_{AV}) \equiv D(1+3\lambda^2)/2|\lambda|$ ; this assumes that  $g_A$  and  $g_V$  are real.

VALUE (°)	CL%	DOCUMENT ID	TECN	COMMENT
<b>180.017 ± 0.026 OUR AVERAGE</b>				
180.012 ± 0.028	68	CHUPP 12	CNTR	Cold <i>n</i> , polarized > 91%
180.04 ± 0.09		SOLDNER 04	CNTR	Cold <i>n</i> , polarized
180.08 ± 0.13		LISING 00	CNTR	Polarized > 93%
• • • We do not use the following data for averages, fits, limits, etc. • • •				
180.013 ± 0.028		MUMM 11	CNTR	See CHUPP 12
179.71 ± 0.39		EROZOLIM... 78	CNTR	Cold <i>n</i> , polarized
180.35 ± 0.43		EROZOLIM... 74	CNTR	Cold <i>n</i> , polarized
181.1 ± 1.3		<sup>1</sup> KROPF 74	RVUE	<i>n</i> decay
180.14 ± 0.22		STEINBERG 74	CNTR	Cold <i>n</i> , polarized

<sup>1</sup> KROPF 74 reviews all data through 1972.

## TRIPLE CORRELATION COEFFICIENT $D$

These are measurements of the component of  $n$  spin perpendicular to the decay plane in  $\beta$  decay. Should be zero if  $T$  invariance is not violated.

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>-1.2 ± 2.0 OUR AVERAGE</b>			
-0.94 ± 1.89 ± 0.97	CHUPP 12	CNTR	Cold <i>n</i> , polarized > 91%
-2.8 ± 6.4 ± 3.0	SOLDNER 04	CNTR	Cold <i>n</i> , polarized
-6 ± 12 ± 5	LISING 00	CNTR	Polarized > 93%
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.96 ± 1.89 ± 1.01	MUMM 11	CNTR	See CHUPP 12
+22 ± 30	EROZOLIM... 78	CNTR	Cold <i>n</i> , polarized
-27 ± 50	<sup>1</sup> EROZOLIM... 74	CNTR	Cold <i>n</i> , polarized
-11 ± 17	STEINBERG 74	CNTR	Cold <i>n</i> , polarized

<sup>1</sup> EROZOLIMSKII 78 says asymmetric proton losses and nonuniform beam polarization may give a systematic error up to  $30 \times 10^{-4}$ , thus increasing the EROZOLIMSKII 74 error to  $50 \times 10^{-4}$ . STEINBERG 74 and STEINBERG 76 estimate these systematic errors to be insignificant in their experiment.

## TRIPLE CORRELATION COEFFICIENT $R$

Another test of time-reversal invariance.  $R$  measures the polarization of the electron in the direction perpendicular to the plane defined by the neutron spin and the electron momentum.  $R = 0$  for  $T$  invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>+0.004 ± 0.012 ± 0.005</b>	<sup>1</sup> KOZELA 12	CNTR	Mott polarimeter
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+0.008 ± 0.015 ± 0.005	KOZELA 09	CNTR	See KOZELA 12

<sup>1</sup> KOZELA 12 also measures the polarization of the electron along the direction of the neutron spin. This is nonzero in the Standard Model; the correlation coefficient is  $N = +0.067 \pm 0.011 \pm 0.004$ .

## *n* REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

MOHR 16	arXiv:1507.07956	P.J. Mohr, D.B. Newell, B.N. Taylor	(NIST)
Accepted for publication in RMP			
ABE 15C	PR D91 072006	K. Abe et al.	(Super-Kamiokande Collab.)
ARZUMANOV 15	PL B745 79	S. Arzumanov et al.	(ILLG, KIAE)
PENDLEBURY 15	PR D92 092003	J.M. Pendlebury et al.	(ETHZ, PSI, SUSS)
SEREBROV 15	PR C92 065501	A.P. Serebrov et al.	(PNPI, ILLG, IOFF)
EPSTEIN 14	PR D90 074027	Z. Epstein, G. Paz, J. Roy	(UMD, WAYN)
MYERS 14	PRL 113 262506	L.S. Myers et al.	(COMPTON/MAX-lab Collab.)
SEREBROV 14	JETPL 99 4	A.P. Serebrov et al.	(PNPI, ILL, IOFF)
MENDENHALL 13	PR C87 032501	M.P. Mendenhall et al.	(UCNA Collab.)
MUND 13	PRL 110 172502	D. Mund et al.	(HEID, ILLG)
YUE 13	PRL 111 222501	A.T. Yue et al.	(UMD, NIST, TENN, ORNL+)
ARZUMANOV 12	JETPL 95 224	S.S. Arzumanov et al.	(KIAE)
Translated from ZETFP 95 248			
CHUPP 12	PR C86 035505	T.E. Chupp et al.	(MICH, UCB, WASH+)
GRIESSHAM... 12	PNP 67 841	H.W. Griesshammer et al.	(GWU, MCHS+)
KOZELA 12	PR C85 045501	A. Kozela et al.	(NTRV Collab.)
MOHR 12	RMP 84 1527	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
PLASTER 12	PR C86 065501	B. Plaster et al.	(UCNA Collab.)
STEYERL 12	PR C85 065503	A. Steyerl et al.	(URI, SUSS)
BRESSI 11	PR A83 052101	G. Bressi et al.	(LEGN, PAVII, PADO, TRST+)
DUBBERS 11	RMP 83 1111	D. Dubbers, M.G. Schmidt	(HEID)
MUMM 11	PRL 107 102301	H.P. Mumm et al.	(NIST, WASH, MICH, LBL+)
WIETFELDT 11	RMP 83 1173	F.E. Wietfeldt, G.L. Greene	(TUL, TENN)
COOPER 10	PR C81 035503	R.L. Cooper et al.	(MICH, NIST, TUL+)

# Baryon Particle Listings

*n*, *N*'s and  $\Delta$ 's

LIU	10	PRL 105 181803	J. Liu <i>et al.</i>	(UCNA Collab.)
Also		PRL 105 219903 (errat.)	J. Liu <i>et al.</i>	(UCNA Collab.)
PICHLMAIER	10	PL B693 221	A. Pichlmaier <i>et al.</i>	(MUNT, PNPI, ILLG)
ALTAREV	09A	PR D80 032003	I. Altarev <i>et al.</i>	(MUNT, RAL, CAEN+)
KOZELA	09	PR L102 172301	A. Kozela <i>et al.</i>	(JAGL, CRAC, PSI, CAEN+)
LAMOREAUX	09	JP G36 104002	S.K. Lamoreaux, R. Golub	(YALE, NCSU)
MOHAPATRA	09	JP G36 104006	R.N. Mohapatra	(UMD)
PATTIE	09	PRL 102 012301	R.W. Pattie Jr. <i>et al.</i>	(Los Alamos UCNA Collab.)
ABELE	08	PNPN 60 1	H. Abele	(HEID)
MOHR	08	RMP 80 633	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
SCHUMANN	08	PRL 100 151801	M. Schumann <i>et al.</i>	(HEID, ILLG, KARL+)
SEREBROV	08	PL B663 181	A.P. Serebrov <i>et al.</i>	(PNPI, IOFF, ILLG+)
BAKER	07	PRL 98 149102	C.A. Baker <i>et al.</i>	(RAL, SUSS, ILLG)
BAN	07	PRL 99 161603	G. Ban <i>et al.</i>	(CAEN, JAGL, PSI, JINR+)
BELUSHKIN	07	PR C75 035202	M.A. Belushkin, H.W. Hammer, U.-G. Meissner	(BONN+)
LAMOREAUX	07	PRL 98 149101	S.K. Lamoreaux, R. Golub	(YALE, NCSU)
SCHUMANN	07	PRL 99 191803	M. Schumann <i>et al.</i>	(HEID, ILLG, KARL+)
SILENKO	07	PNPL 4 468	A.Ya. Silenko	(Belarusian U.)
Also		Translated from PFECAY 6 784		
BAKER	06	PRL 97 131801	C.A. Baker <i>et al.</i>	(RAL, SUSS, ILLG)
BEREZHIANI	06	PRL 96 081801	Z. Bereziani, L. Bento	(Aguila U., LISB)
NICO	06	NAT 444 1059	J.S. Nico <i>et al.</i>	(NIST, TULN, MICH, UMD+)
SEVERIJNS	06	RMP 78 991	N. Severijns, M. Beck, O. Naviliat-Cuncic	(PNPI, IOFF, ILLG+)
KREUZ	05	PL B619 263	M. Kreuz <i>et al.</i>	(HEID, ILLG, MANZ, KARL+)
MOHR	05	RMP 77 1	P.J. Mohr, B.N. Taylor	(NIST)
NICO	05	PR C71 055502	J.S. Nico <i>et al.</i>	(NIST, TULN, IND, TENN+)
SCHUMACHER	05	PNPN 55 567	M. Schumacher	(GOET)
SEREBROV	05	PL B605 72	A. Serebrov <i>et al.</i>	(PNPI, JINR, ILLG)
Also		SPL 48 867	A.P. Serebrov <i>et al.</i>	(PNPI, JINR, ILLG)
Also		Translated from UFN 175 905		
WIETFFELDT	05	MPL A20 1783	F.E. Wietfeldt	(TULN)
SOLDNER	04	PL B581 49	T. Soldner <i>et al.</i>	(ILLG, MUNT)
DEWEY	03	PRL 91 152302	M.S. Dewey <i>et al.</i>	(NIST, TULN, IND+)
KOSSERT	03	EJP A16 259	K. Kossert <i>et al.</i>	(Mainz MAMI Collab.)
Also		PRL 88 162301	K. Kossert <i>et al.</i>	(Mainz MAMI Collab.)
LUNDIN	03	PRL 90 192501	M. Lundin <i>et al.</i>	(PERKEO-II Collab.)
ABELE	02	PRL 89 211801	H. Abele <i>et al.</i>	(LEUV, SUSS, KIAE, PNPI)
BECK	02	JETPL 76 332	M. Beck <i>et al.</i>	
Also		Translated from ZETFP 76 332		
BYRNE	02	JP G28 1325	J. Byrne <i>et al.</i>	
CHUNG	02B	PR D66 032004	J. Chung <i>et al.</i>	(SOUDAN-2 Collab.)
MOSTOVOY	01	PAN 64 1955	Yu.A. Mostovoi <i>et al.</i>	
Also		Translated from YAF 64 2040		
ARZUMANOV	00	PL B483 15	S. Arzumano <i>et al.</i>	
GAL	00	PR C61 028201	A. Gal	
KOLB	00	PRL 85 1388	N.R. Kolb <i>et al.</i>	
LAMOREAUX	00	PR D61 051301	S.K. Lamoreaux, R. Golub	
LEVCHUK	00	NP A674 449	M.I. Levchuk, A.I. L'vov	(BELA, LEBD)
LISING	00	PR C62 055501	L.J. Lising <i>et al.</i>	(NIST emit Collab.)
HARRIS	99	PRL 82 904	P.G. Harris <i>et al.</i>	
KESSLER	99	PL A255 221	E.G. Kessler Jr <i>et al.</i>	
MOHR	99	JPCRD 28 1713	P.J. Mohr, B.N. Taylor	(NIST)
Also		RMP 72 351	P.J. Mohr, B.N. Taylor	(NIST)
SEREBROV	98	JETP 86 1074	A.P. Serebrov <i>et al.</i>	
Also		Translated from ZETFP 113 1963		
ABELE	97D	PL B407 212	H. Abele <i>et al.</i>	(HEIDP, ILLG)
KOPECKY	97	PR C56 2229	S. Kopecky <i>et al.</i>	
LIAUD	97	NP A612 53	P. Liaud <i>et al.</i>	(ILLG, LAPP)
YEROZOLIM...	97	PL B412 240	B.G. Eroozolimsky <i>et al.</i>	(HARV, PNPI, KIAE)
ALTAREV	96	PAN 59 1152	I.S. Altarev <i>et al.</i>	(PNPI)
Also		Translated from YAF 59 1204		
BONDAREN...	96	JETPL 64 416	L.N. Bondarenko <i>et al.</i>	(KIAE)
Also		Translated from ZETFP 64 3832		
BYRNE	96	EPL 33 187	J. Byrne <i>et al.</i>	(SUSS, ILLG)
MOSTOVOY	96	PAN 59 968	Y.A. Mostovoy	(KIAE)
Also		Translated from YAF 59 1013		
NORMAN	96	PR D53 4086	E.B. Norman, J.N. Bahcall, M. Goldhaber	(LBL+)
IGNATOVICH	95	JETPL 62 1	V.K. Ignatovich	(JINR)
Also		Translated from ZETFP 62 3		
KOESTER	95	PR C51 3363	L. Koester <i>et al.</i>	(MUNT, JINR, LATV)
KOPECKY	95	PRL 74 2427	S. Kopecky <i>et al.</i>	
KUZNETSOV	95	PRL 75 794	I.A. Kuznetsov <i>et al.</i>	(PNPI, KIAE, HARV+)
SCHRECK...	95	PL B349 427	K. Schreckenbach <i>et al.</i>	(MUNT, ILLG, LAPP)
BALDO...	94	ZPHY C63 409	M. Baldo-Ceolin <i>et al.</i>	(HEID, ILLG, PADO+)
DIFILIPPO	94	PRL 73 1481	F. Difilippo <i>et al.</i>	(MIT)
Also		PRL 71 1998	V. Natarajan <i>et al.</i>	(MIT)
GOLUB	94	PRPL 237C 1	R. Golub, K. Lamoreaux	(HAHN, WASH)
MAMPE	93	JETPL 57 82	B. Mampe <i>et al.</i>	(KIAE)
Also		Translated from ZETFP 57 77		
ALTAREV	92	PL B276 242	I.S. Altarev <i>et al.</i>	(PNPI)
NESVIZHEV...	92	JETP 75 405	V.V. Nesvizhevsky <i>et al.</i>	(PNPI, JINR)
Also		Translated from ZETFP 102 740		
ALBERICO	91	NP A523 488	W.M. Alberico, A. de Pace, M. Pignone	(TORI)
DUBBERS	91	NP A527 239c	D. Dubbers	(ILLG)
Also		EPL 11 195	D. Dubbers, W. Mampe, J. Dohner	(ILLG, HEID)
EROZOLIM...	91	PL B263 33	B.G. Eroozolimsky <i>et al.</i>	(PNPI, KIAE)
Also		SJNP 52 999	B.G. Eroozolimsky <i>et al.</i>	(PNPI, KIAE)
Also		Translated from YAF 52 1563		
EROZOLIM...	91B	SJNP 53 260	B.G. Eroozolimsky, Y.A. Mostovoy	(KIAE)
Also		Translated from YAF 53 418		
SCHMIEDM...	91	PRL 66 1015	J. Schmiedmayer <i>et al.</i>	(TUW, ORNL)
WOOLCOCK	91	MPL A6 2579	W.S. Woolcock	(CANB)
ALFIMENKOV	90	JETPL 52 373	V.P. Alfimenkov <i>et al.</i>	(PNPI, JINR)
Also		Translated from ZETFP 52 384		
BALDO...	90	PL B236 95	M. Baldo-Ceolin <i>et al.</i>	(PADO, PAVI, HEIDP+)
BERGER	90	PL B240 237	C. Berger <i>et al.</i>	(FREJUS Collab.)
BRESSI	90	NC 103A 731	G. Bressi <i>et al.</i>	(PAVI, ROMA, MILA)
BYRNE	90	PRL 65 289	J. Byrne <i>et al.</i>	(SUSS, NBS, SCOT, CBNM)
GREEN	90	JP G16 L75	K. Green, D. Thompson	(RAL)
RAMSEY	90	ARNPS 40 1	N.F. Ramsey	(HARV)
ROSE	90	PL B234 460	K.W. Rose <i>et al.</i>	(GOET, MPCM, MANZ)
ROSE	90B	NP A514 621	K.W. Rose <i>et al.</i>	(GOET, MPCM)
SMITH	90	PL B234 191	K.F. Smith <i>et al.</i>	(SUSS, RAL, HARV+)
BRESSI	89	ZPHY C43 175	G. Bressi <i>et al.</i>	(INFN, MILA, PAVI, ROMA)
DOVER	89	NIM A284 13	C.B. Dover, A. Gal, J.M. Richard	(BNL, HEBR+)
KOSSAKOW...	89	NP A503 473	R. Kossakowski <i>et al.</i>	(LAPP, SAVO, ISNG+)
MAMPE	89	PRL 63 593	W. Mampe <i>et al.</i>	(ILLG, RISL, SUSS, URI)
MOHAPATRA	89	NIM A284 1	R.N. Mohapatra	(UMD)
PAUL	89	ZPHY C45 25	W. Paul <i>et al.</i>	(BONN, WUPP, MPIH, ILLG)
SCHMIEDM...	89	NIM A284 137	J. Schmiedmayer, H. Rauch, P. Riels	(WIEN)
BAUMANN	88	PR D37 3107	J. Baumann <i>et al.</i>	(BAYR, MUNI, ILLG)
KOESTER	88	ZPHY A329 229	L. Koester, W. Waschkowski, J. Meier	(MUNI, MUNT)
LAST	88	PRL 60 995	I. Last <i>et al.</i>	(HEIDP, ILLG, ANL)
SCHMIEDM...	88	PRL 61 1065	J. Schmiedmayer, H. Rauch, P. Riels	(TUW)
Also		PRL 61 2509 (erratum)	J. Schmiedmayer, H. Rauch, P. Riels	(TUW)
SPIVAK	88	JETP 67 1735	P.E. Spivak	(KIAE)
Also		Translated from ZETFP 94 1		
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RISC, NBS)
ALEKSANDR...	86	SJNP 44 900	Yu.A. Aleksandrov <i>et al.</i>	
Also		Translated from YAF 44 1384		
ALTAREV	86	JETPL 44 460	I.S. Altarev <i>et al.</i>	(PNPI)
Also		Translated from ZETFP 44 360		

BOPP	86	PRL 56 919	P. Bopp <i>et al.</i>	(HEIDP, ANL, ILLG)
Also		ZPHY C37 179	E. Klempt <i>et al.</i>	(HEIDP, ANL, ILLG)
CRESTI	86	PL B177 206	M. Cresti <i>et al.</i>	(PADO)
Also		PL B200 587 (erratum)	M. Cresti <i>et al.</i>	(PADO)
GREENE	86	PRL 56 819	G.L. Greene <i>et al.</i>	(NBS, ILLG)
KOESTER	86	Physica B137 282	L. Koester <i>et al.</i>	
KOSVINTSEV	86	JETPL 44 571	V.Y. Kosvintsev, V.I. Morozov, G.I. Terekhov	(KIAE)
Also		Translated from ZETFP 44 444		
TAKITA	86	PR D34 902	M. Takita <i>et al.</i>	(KEK, TOKY+)
DOVER	85	PR C31 1423	C.B. Dover, A. Gal, J.M. Richard	(BNL)
FIDECARO	85	PL 156B 122	G. Fidencaro <i>et al.</i>	(CERN, ILLG, PADO+)
PARK	85B	NP B252 261	H.S. Park <i>et al.</i>	(IMB Collab.)
BATTISTONI	84	PL 133B 454	G. Battistoni <i>et al.</i>	(NUSEX Collab.)
JONES	84	PRL 52 720	T.W. Jones <i>et al.</i>	(IMB Collab.)
PENDLEBURY	84	PL 136B 327	J.M. Pendlebury <i>et al.</i>	(SUSS, HARV, ILLG+)
CHERRY	83	PRL 50 1354	M.L. Cherry <i>et al.</i>	(PENN, BNL)
DOVER	83	PR D27 1090	C.B. Dover, A. Gal, J.M. Richard	(BNL)
KABIR	83	PRL 51 231	P.K. Kabir	(HARV)
MOSTOVOY	83	JETPL 37 196	Y.A. Mostovoy	(KIAE)
Also		Translated from ZETFP 37 162		
ROY	83	PR D28 1770	A. Roy <i>et al.</i>	(TATA)
VAIDYA	83	PR D27 486	S.C. Vaidya <i>et al.</i>	(TATA)
GAHLER	82	PR D25 2887	R. Gahler, J. Kalus, W. Mampe	(BAYR, ILLG)
GREENE	82	Metrologia 18 93	G.L. Greene <i>et al.</i>	(YALE, HARV, ILLG+)
ALTAREV	81	PL 102B 13	I.S. Altarev <i>et al.</i>	(PNPI)
BARABANOV	80	JETPL 32 359	I.R. Barabanov <i>et al.</i>	(PNPI)
Also		Translated from ZETFP 32 384		
BYRNE	80	PL 92B 274	J. Byrne <i>et al.</i>	(SUSS, RL)
KOSVINTSEV	80	JETPL 31 236	Y.Y. Kosvintsev <i>et al.</i>	(JINR)
Also		Translated from ZETFP 31 257		
MOHAPATRA	80	PRL 44 1316	R.N. Mohapatra, R.E. Marshak	(CUNY, VPI)
ALTAREV	79	JETPL 29 730	I.S. Altarev <i>et al.</i>	(PNPI)
Also		Translated from ZETFP 29 794		
EROZOLIM...	79	SJNP 30 356	B.G. Eroozolimsky <i>et al.</i>	(KIAE)
Also		Translated from YAF 30 692		
NORMAN	79	PRL 43 1226	E.B. Norman, A.G. Seamster	(WASH)
BONDAREN...	78	JETPL 28 303	L.N. Bondarenko <i>et al.</i>	(KIAE)
Also		Translated from ZETFP 28 328		
EROZOLIM...	78	SJNP 28 48	P.G. Eroozolimsky <i>et al.</i>	(KIAE)
Also		Translated from YAF 28 98		
STRATOWA	78	PR D18 3970	C. Stratowa, R. Dobrozemsky, P. Weinzierl	(SEIB)
EROZOLIM...	77	JETPL 23 663	B.G. Eroozolimsky <i>et al.</i>	(KIAE)
Also		Translated from ZETFP 23 720		
KOESTER	76	PRL 36 1021	L. Koester <i>et al.</i>	
STEINBERG	76	PR D13 2469	R.I. Steinberg <i>et al.</i>	(YALE, ISNG)
DOBROZE...	75	PR D11 510	R. Dobrozemsky <i>et al.</i>	(SEIB)
KROHN	75	PL 55B 175	V.E. Krohn, G.R. Ringo	(ANL)
EROZOLIM...	74	JETPL 20 345	B.G. Eroozolimsky <i>et al.</i>	
Also		Translated from ZETFP 20 745		
KROPF	74	ZPHY 267 129	H. Kropf, E. Paul	(LINZ)
Also		NP A154 160	H. Paul	(VIEN)
STEINBERG	74	PRL 33 41	R.I. Steinberg <i>et al.</i>	(YALE, ISNG)
COHEN	73	JPCRD 2 664	E.R. Cohen, B.N. Taylor	(RISC, NBS)
KROHN	73	PR D8 1305	V.E. Krohn, G.R. Ringo	
CHRISTENSEN	72	PR D5 1628	C.J. Christensen <i>et al.</i>	(RIS O)
CHRISTENSEN	70	PR C1 1693	C.J. Christensen, V.E. Krohn, G.R. Ringo	(ANL)
EROZOLIM...	70C	PL 33B 351	B.G. Eroozolimsky <i>et al.</i>	(KIAE)
GRIGOREV	68	SJNP 6 239	V.K. Grigoriev <i>et al.</i>	(ITEP)
Also		Translated from YAF 6 329		
KROHN	66	PR 148 1303	V.E. Krohn, G.R. Ringo	
LEE	56	PR 104 254	T.D. Lee, C.N. Yang	(COLU, BNL)

## N AND Δ RESONANCES

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### I. Introduction

The excited states of the nucleon have been studied in a large number of formation and production experiments. The Breit-Wigner masses and widths, the pole positions, and the elasticities of the *N* and  $\Delta$  resonances in the Baryon Summary Table come largely from partial-wave analyses of  $\pi N$  total, elastic, and charge-exchange scattering data. The most comprehensive analyses were carried out by the Karlsruhe-Helsinki (KH80) [1], Carnegie Mellon-Berkeley (CMB80) [2], and George Washington U (GWU) [3] groups. Partial-wave analyses have also been performed on much smaller  $\pi N$  reaction data sets to get  $\eta N$ ,  $K\Lambda$ , and  $K\Sigma$  branching fractions (see the Listings for references). Other branching fractions come from analyses of  $\pi N \rightarrow \pi\pi N$  data.

In recent years, a large amount of data on photoproduction of many final states has been accumulated, and these data are beginning to tell us much about the properties of baryon resonances. A survey of data on photoproduction can be found in the proceedings of recent conferences [4] and workshops [5], and in recent reviews [6,7].

See key on page 601

# Baryon Particle Listings

## $N$ 's and $\Delta$ 's

Table 1. The status of the  $N$  resonances. Only those with an overall status of \*\*\* or \*\*\*\* are included in the main Baryon Summary Table.

Status as seen in											
Particle	$J^P$	overall	$N\gamma$	$N\pi$	$N\eta$	$N\sigma$	$N\omega$	$\Lambda K$	$\Sigma K$	$N\rho$	$\Delta\pi$
$N$	$1/2^+$	****									
$N(1440)$	$1/2^+$	****	****	****		***				*	***
$N(1520)$	$3/2^-$	****	****	****	***					***	***
$N(1535)$	$1/2^-$	****	****	****	****					**	*
$N(1650)$	$1/2^-$	****	****	****	***			***	**	**	***
$N(1675)$	$5/2^-$	****	****	****	*			*		*	***
$N(1680)$	$5/2^+$	****	****	****	*	**				***	***
$N(1700)$	$3/2^-$	***	**	***	*			*	*	*	***
$N(1710)$	$1/2^+$	****	****	****	***		**	****	**	*	**
$N(1720)$	$3/2^+$	****	****	****	***			**	**	**	*
$N(1860)$	$5/2^+$	**		**						*	*
$N(1875)$	$3/2^-$	***	***	*			**	***	**		***
$N(1880)$	$1/2^+$	**	*	*		**		*			
$N(1895)$	$1/2^-$	**	**	*	**			**	*		
$N(1900)$	$3/2^+$	***	***	**	**		**	***	**	*	**
$N(1990)$	$7/2^+$	**	**	**					*		
$N(2000)$	$5/2^+$	**	**	*	**			**	*	**	
$N(2040)$	$3/2^+$	*		*							
$N(2060)$	$5/2^-$	**	**	**	*				**		
$N(2100)$	$1/2^+$	*		*							
$N(2120)$	$3/2^-$	**	**	**				*	*		
$N(2190)$	$7/2^-$	****	***	****		*	**	**		*	
$N(2220)$	$9/2^+$	****		****							
$N(2250)$	$9/2^-$	****		****							
$N(2300)$	$1/2^+$	**		**							
$N(2570)$	$5/2^-$	**		**							
$N(2600)$	$11/2^-$	***		***							
$N(2700)$	$13/2^+$	**		**							
****	Existence is certain, and properties are at least fairly well explored.										
***	Existence is very likely but further confirmation of decay modes is required.										
**	Evidence of existence is only fair.										
*	Evidence of existence is poor.										

## II. Naming scheme for baryon resonances

In the past, when nearly all resonance information came from elastic  $\pi N$  scattering, it was common to label resonances with the incoming partial wave  $L_{2I,2J}$ , as in  $\Delta(1232)P_{33}$  and  $N(1680)F_{15}$ . However, most recent information has come from  $\gamma N$  experiments. Therefore, we have replaced  $L_{2I,2J}$  with the spin-parity  $J^P$  of the state, as in  $\Delta(1232)3/2^+$  and  $N(1680)5/2^+$ ; this name gives intrinsic properties of the resonance that are independent of the specific particles and reactions used to study them. This applies equally to all baryons, including  $\Xi$  resonances and charm baryons that are not produced in formation experiments. We do not, however, attach the mass or spin-parity to the names of the ground-state (“stable”) baryons  $N, \Lambda, \Sigma, \Xi, \Omega, \Lambda_c, \dots$ .

## III. Using the $N$ and $\Delta$ listings

Tables 1 and 2 list all the  $N$  and  $\Delta$  entries in the Baryon Listings and give our evaluation of the overall status and the status channel by channel. Only the established resonances (overall status 3 or 4 stars) are promoted to the Baryon Summary Table. We long ago omitted from the Listings information from old

analyses, prior to KH80 and CMB80, which can be found in earlier editions. A rather complete survey of older results was given in our 1982 edition [8].

Table 2. The status of the  $\Delta$  resonances. Only those with an overall status of \*\*\* or \*\*\*\* are included in the main Baryon Summary Table.

Status as seen in											
Particle	$J^P$	overall	$N\gamma$	$N\pi$	$N\eta$	$N\sigma$	$N\omega$	$\Lambda K$	$\Sigma K$	$N\rho$	$\Delta\pi$
$\Delta(1232)$	$3/2^+$	****	****	****	F						
$\Delta(1600)$	$3/2^+$	***	***	***	o					*	***
$\Delta(1620)$	$1/2^-$	****	***	****	r					***	***
$\Delta(1700)$	$3/2^-$	****	****	****	b					**	***
$\Delta(1750)$	$1/2^+$	*		*	i						
$\Delta(1900)$	$1/2^-$	**	**	**			d		**	**	**
$\Delta(1905)$	$5/2^+$	****	****	****			d		***	**	**
$\Delta(1910)$	$1/2^+$	****	**	****			e		*	*	**
$\Delta(1920)$	$3/2^+$	***	**	***			n	***		**	**
$\Delta(1930)$	$5/2^-$	***		***							
$\Delta(1940)$	$3/2^-$	**	**	*	F						
$\Delta(1950)$	$7/2^+$	****	****	****	o				***	*	***
$\Delta(2000)$	$5/2^+$	**			r					**	**
$\Delta(2150)$	$1/2^-$	*		*	b						
$\Delta(2200)$	$7/2^-$	*		*	i						
$\Delta(2300)$	$9/2^+$	**	**	**	d						
$\Delta(2350)$	$5/2^-$	*		*	d						
$\Delta(2390)$	$7/2^+$	*		*	e						
$\Delta(2400)$	$9/2^-$	**	**	**	n						
$\Delta(2420)$	$11/2^+$	****	*	****							
$\Delta(2750)$	$13/2^-$	**		**							
$\Delta(2950)$	$15/2^+$	**		**							
****	Existence is certain, and properties are at least fairly well explored.										
***	Existence is very likely but further confirmation of decay modes is required.										
**	Evidence of existence is only fair.										
*	Evidence of existence is poor.										

As a rule, we award an overall status \*\*\*\* or \*\*\* only to those resonances which are derived from analyses of data sets that include precision differential cross sections and polarization observables, and are confirmed by independent analyses. All other signals are given \*\* or \* status. We do not consider new results that are not accompanied by proper error evaluation. The following criteria are guidelines for future error analysis.

1. Uncertainties in resonance parameters: The publication must have a detailed discussion on how the uncertainties of parameters were estimated and why the author(s) believe that they approximately represent real uncertainties. This requires that the error estimates go beyond the simple fit error as e.g. given by MINUIT, and the robustness of the results must be demonstrated.

2. Fit quality: Concrete measures for the fit quality must be provided. The reduced global  $\chi^2$  value of the fit, while useful, is insufficient. Other possibilities include quoting variations of local  $\chi^2$  value in kinematic regions where evidence for new states or significantly improved information on resonance parameters is claimed.

3. Weight factors in observables: Analyses often use weight factors for certain data sets to either increase or reduce their

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## $N$ 's and $\Delta$ 's

impact on the results. This has been particularly important when polarization observables are involved, which usually are very sensitive to amplitude interferences but often have much poorer statistics than differential cross section data. To evaluate sensitivities, the resulting resonance parameters should be checked against variations of the specific weight factors.

Claims of evidence for new baryon states must be based on a sufficiently complete set of partial waves in the fit. The robustness of signals must be demonstrated, e.g. by examining the effect of higher waves in the fit.

### IV. Properties of resonances

Resonances are defined by poles of the  $S$ -matrix, whether in scattering, production or decay matrix elements. These are poles in the complex plane in  $s$ , as discussed in the new review on *Resonances*. As traditional we quote here the pole positions in the complex energy  $w = \sqrt{s}$  plane. Crucially, the position of the pole of the  $S$ -matrix is independent of the process, and the production and decay properties factorize. This is the rationale for listing the pole position first for each resonance.

These key properties of the  $S$ -matrix pole are in contrast to other quantities related to resonance phenomena, such as Breit-Wigner parameters or any  $K$ -matrix pole. Thus, Breit-Wigner parameters depend on the formalism used such as angular-momentum barrier factors, or cut-off parameters, and the assumed or modeled background. However, the accurate determination of pole parameters from the analysis of data on the real energy axis is not necessarily simple, or even straightforward. It requires the implementation of the correct analytic structure of the relevant (often coupled) channels. The example in the meson sector of the  $\sigma$ -pole highlights the need to incorporate right and left hand cut analyticity (and their relation imposed by crossing symmetry) into a dispersive analysis to obtain a robust determination of the pole position, for a very short-lived state close to the lowest threshold. The development of general methods that are simpler to implement in the baryon sector is a research problem of current interest, often exploiting techniques introduced long ago when the experimental data were far poorer than those presently available for reactions like  $\gamma N \rightarrow \pi N$  [9]. No consensus yet exists for the use of any particular method, beyond the need to incorporate the general properties mentioned here and discussed more fully in the review of *Resonances*.

The tables will change in the following way: Pole positions, elastic and normalized inelastic residues appear first, followed by the Breit-Wigner mass, width, decay modes and branching ratios. Finally, we give the photon decay amplitudes first as complex values at the pole position and then as real numbers at the Breit-Wigner mass.

### V. Photoproduction

A new approach to the nucleon excitation spectrum is provided by dedicated facilities at the Universities of Bonn, Grenoble, and Mainz, and at the national laboratories Jefferson

Lab in the US and SPring-8 in Japan. High-precision cross sections and polarization observables for the photoproduction of pseudoscalar mesons provide a data set that is nearly a “complete experiment,” one that fully constrains the four complex amplitudes describing the spin-structure of the reaction [11]. A large number of photoproduction reactions has been studied.

In pseudo-scalar meson photoproduction, the four independent helicity amplitudes can be expressed in terms of the four CGLN [12] amplitudes allowed by Lorentz and gauge invariance. These amplitudes can be expanded in a series of electric and magnetic multipoles. Except for  $J=1/2$ , one electric and one magnetic multipole contributes to each  $J^P$  combination.

For a given state, these two amplitudes determine the resonance photo-decay helicity amplitudes  $A_{1/2}$  and  $A_{3/2}$ . As described below, this resonance extraction has been carried out either assuming a Breit-Wigner resonance or at the pole.

If a Breit-Wigner parametrization is used, the  $N\gamma$  partial width,  $\Gamma_\gamma$ , is given in terms of the helicity amplitudes  $A_{1/2}$  and  $A_{3/2}$  by

$$\Gamma_\gamma = \frac{k_{\text{BW}}^2}{\pi} \frac{2m_N}{(2J+1)m_{\text{BW}}} \left( |A_{1/2}|^2 + |A_{3/2}|^2 \right). \quad (1)$$

Here  $m_N$  and  $m_{\text{BW}}$  are the nucleon and resonance masses,  $J$  is the resonance spin, and  $k_{\text{BW}}$  is the photon c.m. decay momentum. Most earlier analyses have quoted these real quantities  $A_{1/2}$  and  $A_{3/2}$ .

More recent studies have quoted related complex quantities, evaluated at the T-matrix pole. These complex helicity amplitudes,  $\tilde{A}_{1/2}$  and  $\tilde{A}_{3/2}$ , can be cast onto the form

$$\tilde{A}_h = \sqrt{\frac{\pi(2J+1)w_{\text{pole}}}{m_N k_{\text{pole}}^2}} \frac{\text{Res}(T_h(\gamma N \rightarrow N b))}{\sqrt{\text{Res}(T(N b \rightarrow N b))}} \quad (2)$$

where the residues ( $\text{Res}$ ) are evaluated at the pole position,  $w_{\text{pole}}$ , and  $k_{\text{pole}}^2 = (w_{\text{pole}}^2 - m_N^2)/4w_{\text{pole}}^2$  [13]. For Breit-Wigner amplitudes,  $w_{\text{pole}} = m_{\text{BW}}$  and  $\tilde{A}_h = A_h$ . Similar relations for the photo- and electrocouplings at the pole position can be found in [14,15].

The determination of eight real numbers from four complex amplitudes (with one overall phase undetermined) requires at least seven independent measurements. At least one further measurement is required to resolve discrete ambiguities that result from the fact that data are proportional to squared amplitudes. Photon beams and nucleon targets can be polarized (with linear or circular polarization  $P_\perp$ ,  $P_\odot$  and  $\vec{T}$ , respectively); the recoil polarization of the outgoing baryon  $\vec{R}$  can be measured. The experiments can be divided into three classes: (1) the beam and target are polarized (BT); (2) the beam is polarized and the recoil baryon polarization is measured (BR); (3) the target is polarized and the recoil polarization is measured (TR). Different sign conventions are used in the literature, as summarized in [16].

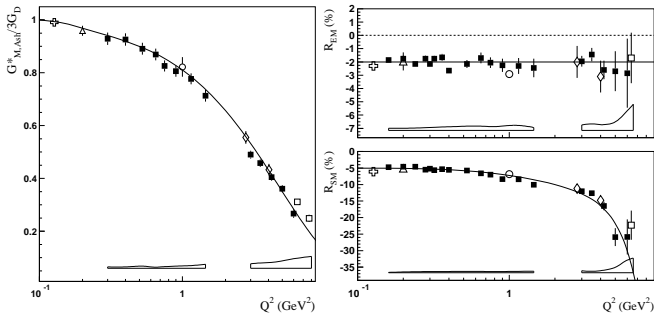


One of the best studied reactions is  $\gamma p \rightarrow \Lambda K^+$ . Published data include differential cross sections, the beam asymmetry  $\Sigma$ , the target asymmetry  $T$ , the recoil polarization  $P$ , and the BR double-polarization variables  $C'_x, C'_z, O'_x$ , and  $O'_z$ . For the photoproduction of pions and etas, off proton and neutron targets, differential cross sections, single- and double-polarization asymmetries have been measured, mainly for pions.

## VI. Electroproduction

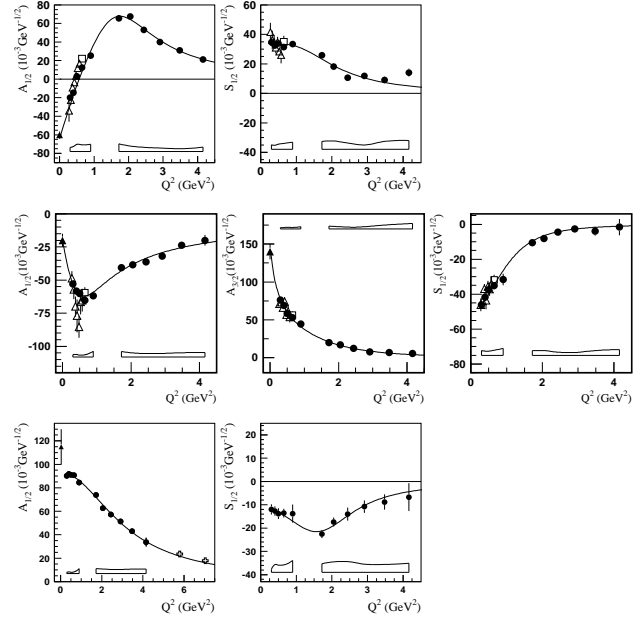
Electro-production of mesons provides information on the internal structure of resonances. The helicity amplitudes are functions of the (squared) momentum transfer  $Q^2 = -(e - e')^2$ , where  $e$  and  $e'$  are the 4-momenta of the incident and scattered electron, and a third amplitude,  $S_{1/2}$ , measures the resonance response to the longitudinal component of the virtual photon. Most data stem from the reactions  $e^- p \rightarrow e^- n \pi^+$  and  $e^- p \rightarrow e^- p \pi^0$  but also the reactions  $e^- p \rightarrow e^- p \eta$ ,  $e^- p \rightarrow e^- p \pi^+ \pi^-$ , and  $e^- p \rightarrow e^- \Lambda(\Sigma^0) K^+$  have been studied. The data and their interpretation are reviewed in Refs. 18,19.

The transition to the  $\Delta(1232)3/2^+$  is often quantified in terms of the magnetic dipole transition moment  $M_{1+}$  (or the magnetic transition form factor  $G_{M,Ash}^*(Q^2)$ ) [20], and the electric and scalar quadrupole transition moments  $E_{1+}$  and  $S_{1+}$ . Figure 1 shows the strength of the  $p \rightarrow \Delta^+$  transition plotted versus the photon virtuality  $Q^2$ . At  $Q^2 = 0$ ,  $M_{1+}$  dominates the resonance transition strength. The two amplitudes  $E_{1+}$  and  $S_{1+}$  imply a quadrupole deformation of the transition to the lowest excited state. The magnitude of  $R_{EM} = E_{1+}/M_{1+}$  remains nearly constant, while the magnitude of  $R_{SM} = S_{1+}/M_{1+}$  increases rapidly up to 25% at the highest  $Q^2$  value. Dynamical models assign most of the quadrupole strength in the  $p \Delta^+$  transition to the effect of a meson cloud around the bare  $\Delta$  states.



**Figure 1:** Left: The magnetic transition form factor for the  $\gamma^* p \rightarrow \Delta^+(1232)$  transition versus the photon virtuality  $Q^2$ . Right: The electric and scalar quadrupole ratios  $R_{EM}$  and  $R_{SM}$ . The different symbols are results from different experiments at JLab (squares, diamonds, circle) and MAMI (triangle, cross). The boxes near the horizontal axis indicate model uncertainties of the squares. Curves to guide the eyes.

Figure 2 shows the transverse and scalar helicity amplitudes for the  $N(1440)1/2^+$ ,  $N(1520)3/2^-$ , and  $N(1535)1/2^-$  resonances from JLab [18]. Similar results have been achieved at Mainz [19]. For the states  $N(1440)1/2^+$  and  $N(1520)3/2^-$ , helicity amplitudes and  $\pi\Delta$  and  $\rho p$  decays were determined at JLab in an analysis of  $\pi^+ \pi^- p$  electroproduction [21]. The data show distinctly different  $Q^2$  dependencies that indicate different internal structures.



**Figure 2:** Transverse and scalar (longitudinal) helicity amplitudes for  $\gamma p \rightarrow N(1440)1/2^+$  (top),  $\gamma p \rightarrow N(1520)3/2^-$  (center), and  $\gamma p \rightarrow N(1535)1/2^-$  (bottom) as extracted from the JLab/CLAS data in  $n\pi^+$  production (full circles), in  $\pi\pi^+\pi^-$  (open triangles), combined single and double pion production (open squares). The solid triangle is the PDG 2014 value at  $Q^2 = 0$ . The open boxes are the model uncertainties of the full circles.

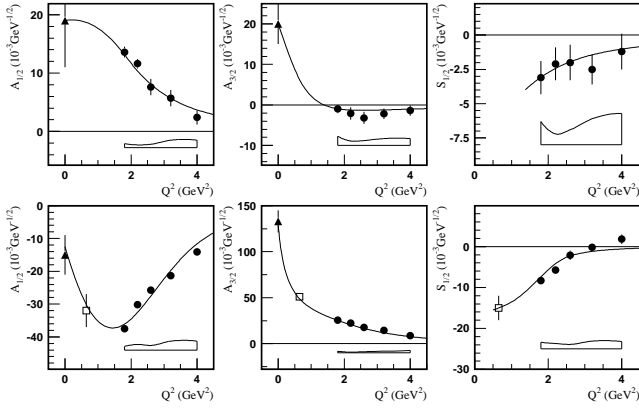
The  $N(1520)3/2^-$  helicity amplitudes reveal the dominance of its three-quark nature: the  $A_{3/2}$  amplitude is large at the photon point and decreases rapidly  $\sim Q^{-5}$  with increasing  $Q^2$ ;  $A_{1/2}$  is small at the photon point, increases rapidly with  $Q^2$  and then falls off with  $\sim Q^{-3}$ . Quantitative agreement with the data is, however, achieved only when meson cloud effects are included.

At high  $Q^2$ , both amplitudes for  $N(1440)1/2^+$  are qualitatively described by light front quark models [22]: at short distances the resonance behaves as expected from a radial excitation of the nucleon. On the other hand,  $A_{1/2}$  changes sign at about  $0.6 \text{ GeV}^2$ . This remarkable behavior has not been observed before for any nucleon form factor or transition amplitude. Obviously, an important change in the structure occurs when the resonance is probed as a function of  $Q^2$ .

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The  $Q^2$  dependence of  $A_{1/2}$  of the  $N(1535)1/2^-$  resonance exhibits the expected  $\sim Q^{-3}$  dependence, except for small  $Q^2$  values where meson cloud effects set in.



**Figure 3:** Transverse and scalar helicity amplitudes for  $\gamma p \rightarrow N(1675)5/2^-$  (top), and  $\gamma p \rightarrow N(1680)5/2^+$  (bottom) as extracted from the JLab/CLAS data in  $n\pi^+$  production (full circles), combined single and double pion production (open square). The solid triangle is the 2014 PDG value at  $Q^2 = 0$ . The open boxes are the model uncertainties of the full circles. The curves are to guide the eye.

Figure 3 shows the transverse and scalar amplitudes for two states in the 3rd nucleon resonance region around 1700 MeV, the  $N(1675)5/2^-$  and  $N(1680)5/2^+$ . These states have nearly degenerate masses and are parity partners. In the quark model picture, the transverse amplitudes for  $N(1675)5/2^-$  on the proton are suppressed due to the Moorhouse selection rule, allowing for a quantitative evaluation of the meson-baryon contributions. The data show significant meson-baryon strength in the  $A_{1/2}$  amplitude even at quite high  $Q^2$ , while  $A_{3/2}$  drops much faster with  $Q^2$ .  $N(1680)5/2^+$  shows qualitatively the features predicted in constituent quark models, a dominant  $A_{3/2}$  at the real photon point that drops rapidly with increasing  $Q^2$ , while  $A_{1/2}$  becomes the dominant contribution at high  $Q^2$ , indicating a switch of the helicity structure in the resonance transition at short distances.

### VII. Partial wave analyses

Several PWA groups are now actively involved in the analysis of the new data. The GWU group maintains a nearly complete database covering reactions from  $\pi N$  and  $KN$  elastic scattering to  $\gamma N \rightarrow N\pi$ ,  $N\eta$ , and  $N\eta'$ . It is presently the only group determining  $\pi N$  elastic amplitudes from scattering data in sliced energy bins. Given the high-precision of photoproduction data already or soon to be collected, the spectrum of  $N$  and  $\Delta$  resonances will in the near future be better known.

Fits to the data are performed by various groups with the aim to understand the reaction dynamics and to identify  $N$  and  $\Delta$  resonances. For practical reasons, approximations have to be made. We mention several analyses here: (1) The Mainz

unitary isobar model [23] focusses on the correct treatment of the low-energy domain. Resonances are added to the unitary amplitude as a sum of Breit-Wigner amplitudes. This model also obtains resonance transition form factors and helicity amplitudes from electroproduction [19]. (2) For  $N\pi$  electroproduction, the Yerevan/JLab group uses both the unitary isobar model and the dispersion relation approach developed in [22]. A phenomenological model was developed to extract resonance couplings and partial decay widths from exclusive  $\pi^+\pi^-p$  electroproduction [21]. (3) Multichannel analyses using K-matrix parameterizations derive background terms from a chiral Lagrangian - providing a microscopical description of the background - (Giessen [24,25]) or from phenomenology (KSU [26,27], Bonn-Gatchina [28]). (4.) Several groups (EBAC-Jlab [29,30], ANL-Osaka [31], Dubna-Mainz-Taipeh [32], Bonn-Jülich [33,34,35], Valencia [36]) use dynamical reaction models, driven by chiral Lagrangians, which take dispersive parts of intermediate states into account. Several other groups have made important contributions. The Giessen group pioneered multichannel analyses of large data sets on pion- and photo-induced reactions [24,25]. The Bonn-Gatchina group included recent high-statistics data and reported systematic searches for new baryon resonances in all relevant partial waves. A summary of their results can be found in Ref. 28.

### References

1. G. Höhler, Pion-Nucleon Scattering, Landolt-Börnstein Vol. I/9b2 (1983), ed. H. Schopper, Springer Verlag.
2. R.E. Cutkosky *et al.*, Baryon 1980, *IV International Conference on Baryon Resonances*, Toronto, ed. N. Isgur, p. 19.
3. R.A. Arndt *et al.*, Phys. Rev. **C74**, 045205 (2006).
4. "Hadron 2011: 14th International Conference on Hadron Spectroscopy", München, Germany, June, 13 - 17, 2011, published in eConf.
5. "NSTAR 2013: 9th International Workshop on the Physics of Excited Nucleons", 27-30 May 2013, Peñíscola, Spain.
6. E. Klempt and J.M. Richard, Rev. Mod. Phys. **82**, 1095 (2010).
7. V. Credé and W. Roberts, Rept. on Prog. in Phys. **76**, 076301 (2013).
8. M. Roos *et al.*, Phys. Lett. **B111**, 1 (1982).
9. A. Svarc *et al.*, Phys. Rev. **C88**, 035206 (2013).
10. R.H. Dalitz, R.G. Moorhouse, Proc. Royal Soc. London **A318**, 279 (1970).
11. C. G. Fasano *et al.*, Phys. Rev. **C46**, 2430 (1992).
12. G.F. Chew *et al.*, Phys. Rev. **106**, 1345 (1957).
13. R.L. Workman, L. Tiator, and A. Sarantsev, Phys. Rev. **C87**, 068201 (2013).
14. N. Suzuki, T. Sato, and T.-S.H. Lee, Phys. Rev. **C82**, 045206 (2010).
15. H. Kamano, Phys. Rev. **C88**, 045203 (2013).
16. A.M. Sandorfi *et al.*, AIP Conf. Proc. **1432**, 219 (2012).
17. R. Beck and A. Thiel, J. Phys. Conf. Ser. **295**, 012023 (2011).

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## $N$ 's and $\Delta$ 's, $N(1440)$

18. I. G. Aznauryan and V. D. Burkert, Prog. in Part. Nucl. Phys. **67**, 1 (2012).
19. L. Tiator *et al.*, Eur. Phys. J. ST **198**, 141 (2011).
20. W.W. Ash, Phys. Lett. **B24**, 165 (1967).
21. V.I. Mokeev *et al.* [CLAS Collab.], Phys. Rev. **C86**, 035203 (2012).
22. I.G. Aznauryan, Phys. Rev. **C67**, 015209 (2003).
23. D. Drechsel, S.S. Kamalov, and L. Tiator, Eur. Phys. J. **A34**, 69 (2007).
24. G. Penner and U. Mosel, Phys. Rev. **C66**, 055211 (2002).
25. G. Penner and U. Mosel, Phys. Rev. **C66**, 055212 (2002).
26. D.M. Manley and E.M. Saleski, Phys. Rev. **D45**, 4002 (1992).
27. M. Shrestha and D.M. Manley, Phys. Rev. **C86**, 055203 (2012).
28. A.V. Anisovich *et al.*, Eur. Phys. J. **A48**, 15 (2012).
29. A. Matsuyama, T. Sato, and T.-S.H. Lee, Phys. Reports **439**, 193 (2007).
30. T. Sato and T.-S.H. Lee, J. Phys. **G36**, 073001 (2009).
31. H. Kamano *et al.*, Phys. Rev. **C88**, 035209 (2013).
32. G.Y. Chen *et al.*, Phys. Rev. **C76**, 035206 (2007).
33. M. Döring *et al.*, Phys. Lett. **B681**, 26 (2009).
34. M. Döring *et al.*, Nucl. Phys. **A829**, 170 (2009).
35. D. Röschner *et al.*, Eur. Phys. J. **A49**, 44 (2013).
36. S. Sarkar, E. Oset, and M.J. Vicente Vacas, Nucl. Phys. **A750**, 294 (2005) [Erratum-ibid. **A780**, 78 (2006)].

$N(1440) \ 1/2^+$

$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \*\*\*\*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

N(1440) POLE POSITION			
REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1369 ± 3	SOKHOYAN	15A	DPWA Multichannel
1363 ± 2 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1359	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1385	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1375 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1386	SHKLYAR	13	DPWA Multichannel
1370 ± 4	ANISOVICH	12A	DPWA Multichannel
1370	SHRESTHA	12A	DPWA $\pi N \rightarrow \pi N$
1363 ± 11	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1383	VRANA	00	DPWA Multichannel
-2xIMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
189 ± 5	SOKHOYAN	15A	DPWA Multichannel
180 ± 4 ± 5	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
162	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
164	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
180 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
277	SHKLYAR	13	DPWA Multichannel
190 ± 7	ANISOVICH	12A	DPWA Multichannel
214	SHRESTHA	12A	DPWA Multichannel
151 ± 13	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
316	VRANA	00	DPWA Multichannel

N(1440) ELASTIC POLE RESIDUE			
MODULUS  r			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
40 to 52 (≈ 46) OUR ESTIMATE			
49 ± 3	SOKHOYAN	15A	DPWA Multichannel
50 ± 1 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
38	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
40	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
52 ± 5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •			
126	SHKLYAR	13	DPWA Multichannel
48 ± 3	ANISOVICH	12A	DPWA Multichannel
44	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
PHASE $\theta$			
VALUE (°)	DOCUMENT ID	TECN	COMMENT
- 82 ± 5	SOKHOYAN	15A	DPWA Multichannel
- 88 ± 1 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
- 98	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
-100 ± 35	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
- 60	SHKLYAR	13	DPWA Multichannel
- 78 ± 4	ANISOVICH	12A	DPWA Multichannel
- 88	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

N(1440) INELASTIC POLE RESIDUE			
The "normalized residue" is the residue divided by $\Gamma_{pole}/2$ .			
Normalized residue in $N\pi \rightarrow N(1440) \rightarrow \Delta\pi, P\text{-wave}$			
MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN COMMENT
27 ± 2	38 ± 5	SOKHOYAN	15A DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
27 ± 2	40 ± 5	ANISOVICH	12A DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(1440) \rightarrow N(\pi\pi)_{S\text{-wave}}^{I=0}$			
MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN COMMENT
21 ± 4	-136 ± 4	SOKHOYAN	15A DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
21 ± 5	-135 ± 7	ANISOVICH	12A DPWA Multichannel

N(1440) BREIT-WIGNER MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1410 to 1450 (≈ 1430) OUR ESTIMATE			
1430 ± 10	SOKHOYAN	15A	DPWA Multichannel
1515 ± 15	SHKLYAR	13	DPWA Multichannel
1485.0 ± 1.2	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1440 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1410 ± 12	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1430 ± 8	ANISOVICH	12A	DPWA Multichannel
1412 ± 2	SHRESTHA	12A	DPWA Multichannel
1439 ± 19	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1518 ± 5	PENNER	02C	DPWA Multichannel
1479 ± 80	VRANA	00	DPWA Multichannel

N(1440) BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
250 to 450 (≈ 350) OUR ESTIMATE			
360 ± 30	SOKHOYAN	15A	DPWA Multichannel
605 ± 90	SHKLYAR	13	DPWA Multichannel
284 ± 18	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
340 ± 70	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
135 ± 10	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
365 ± 35	ANISOVICH	12A	DPWA Multichannel
248 ± 5	SHRESTHA	12A	DPWA Multichannel
437 ± 141	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
668 ± 41	PENNER	02C	DPWA Multichannel
490 ± 120	VRANA	00	DPWA Multichannel

N(1440) DECAY MODES		
The following branching fractions are our estimates, not fits or averages.		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \ N\pi$	55-75 %	
$\Gamma_2 \ N\eta$	<1 %	
$\Gamma_3 \ N\pi\pi$	25-50 %	
$\Gamma_4 \ \Delta(1232)\pi$	20-30 %	
$\Gamma_5 \ \Delta(1232)\pi, P\text{-wave}$	13-27 %	
$\Gamma_6 \ N\sigma$	11-23 %	
$\Gamma_7 \ p\gamma, \text{ helicity}=1/2$	0.035-0.048 %	
$\Gamma_8 \ n\gamma, \text{ helicity}=1/2$	0.02-0.04 %	

# Baryon Particle Listings

## $N(1440)$ , $N(1520)$

### $N(1440)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
63 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel	
56 $\pm$ 2	SHKLYAR	13	DPWA Multichannel	
78.7 $\pm$ 1.6	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$	
68 $\pm$ 4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
51 $\pm$ 5	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
62 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel	
64.8 $\pm$ 0.9	SHRESTHA	12A	DPWA Multichannel	
62 $\pm$ 4	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
57 $\pm$ 1	PENNER	02C	DPWA Multichannel	
72 $\pm$ 5	VRANA	00	DPWA Multichannel	

$\Gamma(N\eta)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0 $\pm$ 1	VRANA	00	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
20 $\pm$ 7	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
21 $\pm$ 8	ANISOVICH	12A	DPWA Multichannel	
6.5 $\pm$ 0.8	SHRESTHA	12A	DPWA Multichannel	
16 $\pm$ 1	VRANA	00	DPWA Multichannel	

$\Gamma(N\sigma)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
17 $\pm$ 6	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17 $\pm$ 7	ANISOVICH	12A	DPWA Multichannel	
27 $\pm$ 1	SHRESTHA	12A	DPWA Multichannel	
12 $\pm$ 1	VRANA	00	DPWA Multichannel	

### $N(1440)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $N(1440) \rightarrow \rho\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS (GeV $^{-1/2}$ )	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
-0.044 $\pm$ 0.005	-40 $\pm$ 8	SOKHOYAN	15A	DPWA Multichannel

### $N(1440)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $N(1440) \rightarrow \rho\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
-0.060 $\pm$ 0.004 OUR ESTIMATE			
-0.061 $\pm$ 0.006	SOKHOYAN	15A	DPWA Multichannel
-0.056 $\pm$ 0.001	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
-0.051 $\pm$ 0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.085 $\pm$ 0.003	SHKLYAR	13	DPWA Multichannel
-0.061 $\pm$ 0.008	ANISOVICH	12A	DPWA Multichannel
-0.084 $\pm$ 0.003	SHRESTHA	12A	DPWA Multichannel
-0.061	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
-0.087	PENNER	02D	DPWA Multichannel

#### $N(1440) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
+0.040 $\pm$ 0.010 OUR ESTIMATE			
0.043 $\pm$ 0.012	ANISOVICH	13B	DPWA Multichannel
0.048 $\pm$ 0.004	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.040 $\pm$ 0.005	SHRESTHA	12A	DPWA Multichannel
0.054	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
0.121	PENNER	02D	DPWA Multichannel

### $N(1440)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $N(1440)$ REFERENCES

For early references, see Physics Letters **111B 1** (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	13B	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	12A	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
CHEN	12A	PR C86 015206	W. Chen <i>et al.</i>	(DUKE, GWU, MSST, ITEP+)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)

BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER	07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohlner	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohlner <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

## $N(1520) 3/2^-$

$$I(J^P) = \frac{1}{2}(\frac{3}{2}^-) \text{ Status: } ***$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $N(1520)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1505 to 1515 ( $\approx$ 1510) OUR ESTIMATE			
1507 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel
1506 $\pm$ 1 $\pm$ 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1515	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1510	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
1510 $\pm$ 5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1492	SHKLYAR	13	DPWA Multichannel
1507 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel
1501	SHRESTHA	12A	DPWA Multichannel
1506 $\pm$ 9	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1504	VRANA	00	DPWA Multichannel

#### -2xIMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
105 to 120 ( $\approx$ 110) OUR ESTIMATE			
111 $\pm$ 3	SOKHOYAN	15A	DPWA Multichannel
115 $\pm$ 2 $\pm$ 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
113	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
120	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
114 $\pm$ 10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
94	SHKLYAR	13	DPWA Multichannel
111 $\pm$ 5	ANISOVICH	12A	DPWA Multichannel
112	SHRESTHA	12A	DPWA Multichannel
122 $\pm$ 9	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
112	VRANA	00	DPWA Multichannel

### $N(1520)$ ELASTIC POLE RESIDUE

#### MODULUS $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
35 $\pm$ 3 OUR ESTIMATE			
36 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel
33 $\pm$ 1 $\pm$ 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
38	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
32	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
35 $\pm$ 2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
27	SHKLYAR	13	DPWA Multichannel
36 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel
35	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

#### PHASE $\theta$

VALUE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
-10 $\pm$ 5 OUR ESTIMATE			
-14 $\pm$ 3	SOKHOYAN	15A	DPWA Multichannel
-15 $\pm$ 1 $\pm$ 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
- 5	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
- 8	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
-12 $\pm$ 5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-35	SHKLYAR	13	DPWA Multichannel
-14 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel
- 7	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

### $N(1520)$ INELASTIC POLE RESIDUE

The "normalized residue" is the residue divided by  $\Gamma_{\text{pole}}/2$ .

#### Normalized residue in $N\pi \rightarrow N(1520) \rightarrow \Delta\pi, S\text{-wave}$

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
0.33 $\pm$ 0.04	155 $\pm$ 15	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.33 $\pm$ 0.05	150 $\pm$ 20	ANISOVICH	12A	DPWA Multichannel

See key on page 601

# Baryon Particle Listings

## $N(1520)$

### Normalized residue in $N\pi \rightarrow N(1520) \rightarrow \Delta\pi$ , $D$ -wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.25 ± 0.03	105 ± 18	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.25 ± 0.03	100 ± 20	ANISOVICH	12A	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow N(1520) \rightarrow N\sigma$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.08 ± 0.03	−45 ± 25	SOKHOYAN	15A	DPWA Multichannel

### $N(1520)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1510 to 1520 (≈ 1515) OUR ESTIMATE</b>			
1516 ± 2	SOKHOYAN	15A	DPWA Multichannel
1505 ± 4	SHKLYAR	13	DPWA Multichannel
1514.5 ± 0.2	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1525 ± 10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1519 ± 4	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1517 ± 3	ANISOVICH	12A	DPWA Multichannel
1512.6 ± 0.5	SHRESTHA	12A	DPWA Multichannel
1522 ± 8	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1509 ± 1	PENNER	02C	DPWA Multichannel
1518 ± 3	VRANA	00	DPWA Multichannel

### $N(1520)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>100 to 125 (≈ 115) OUR ESTIMATE</b>			
113 ± 4	SOKHOYAN	15A	DPWA Multichannel
100 ± 2	SHKLYAR	13	DPWA Multichannel
103.6 ± 0.4	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
120 ± 15	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
114 ± 7	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
114 ± 5	ANISOVICH	12A	DPWA Multichannel
117 ± 1	SHRESTHA	12A	DPWA Multichannel
132 ± 11	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
100 ± 2	PENNER	02C	DPWA Multichannel
124 ± 4	VRANA	00	DPWA Multichannel

### $N(1520)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	55–65 %
$\Gamma_2$ $N\eta$	< 1 %
$\Gamma_3$ $N\pi\pi$	25–35 %
$\Gamma_4$ $\Delta(1232)\pi$	22–34 %
$\Gamma_5$ $\Delta(1232)\pi$ , $S$ -wave	15–23 %
$\Gamma_6$ $\Delta(1232)\pi$ , $D$ -wave	7–11 %
$\Gamma_7$ $N\sigma$	< 2 %
$\Gamma_8$ $p\gamma$	0.31–0.52 %
$\Gamma_9$ $p\gamma$ , helicity=1/2	0.01–0.02 %
$\Gamma_{10}$ $p\gamma$ , helicity=3/2	0.30–0.50 %
$\Gamma_{11}$ $n\gamma$	0.30–0.53 %
$\Gamma_{12}$ $n\gamma$ , helicity=1/2	0.04–0.10 %
$\Gamma_{13}$ $n\gamma$ , helicity=3/2	0.25–0.45 %

### $N(1520)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
<b>55 to 65 OUR ESTIMATE</b>			
61 ± 2	SOKHOYAN	15A	DPWA Multichannel
57 ± 2	SHKLYAR	13	DPWA Multichannel
63.2 ± 0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
58 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
54 ± 3	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
62 ± 3	ANISOVICH	12A	DPWA Multichannel
62.7 ± 0.5	SHRESTHA	12A	DPWA Multichannel
55 ± 5	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
56 ± 1	PENNER	02C	DPWA Multichannel
63 ± 2	VRANA	00	DPWA Multichannel

### $\Gamma(N\eta)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
0 ± 1	SHKLYAR	13	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.1 ± 0.1	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
0.2 ± 0.1	THOMA	08	DPWA Multichannel
0.08 to 0.12	ARNDT	05	DPWA Multichannel
0.23 ± 0.04	PENNER	02C	DPWA Multichannel
0 ± 1	VRANA	00	DPWA Multichannel
0.08 ± 0.01	TIATOR	99	DPWA $\gamma p \rightarrow p\eta$

### $\Gamma(\Delta(1232)\pi, S\text{-wave})/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
19 ± 4	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
19 ± 4	ANISOVICH	12A	DPWA Multichannel
9.3 ± 0.7	SHRESTHA	12A	DPWA Multichannel
15 ± 2	VRANA	00	DPWA Multichannel

### $\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
9 ± 2	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
9 ± 2	ANISOVICH	12A	DPWA Multichannel
6.3 ± 0.5	SHRESTHA	12A	DPWA Multichannel
11 ± 2	VRANA	00	DPWA Multichannel

### $\Gamma(N\sigma)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<2	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<1	SHRESTHA	12A	DPWA Multichannel
<4	THOMA	08	DPWA Multichannel
1 ± 1	VRANA	00	DPWA Multichannel

### $N(1520)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $N(1520) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
−0.023 ± 0.004	−6 ± 5	SOKHOYAN	15A	DPWA Multichannel

#### $N(1520) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$

MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.131 ± 0.006	4 ± 4	SOKHOYAN	15A	DPWA Multichannel

### $N(1520)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $N(1520) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>−0.020 ± 0.005 OUR ESTIMATE</b>			
−0.024 ± 0.004	SOKHOYAN	15A	DPWA Multichannel
−0.019 ± 0.002	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
−0.028 ± 0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
−0.038 ± 0.003	AHRENS	02	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.015 ± 0.001	SHKLYAR	13	DPWA Multichannel
−0.022 ± 0.004	ANISOVICH	12A	DPWA Multichannel
−0.034 ± 0.001	SHRESTHA	12A	DPWA Multichannel
−0.027	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
−0.003	PENNER	02D	DPWA Multichannel
−0.052 ± 0.010 ± 0.007	<sup>2</sup> MUKHOPAD... 98		$\gamma p \rightarrow \eta p$

#### $N(1520) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>0.140 ± 0.010 OUR ESTIMATE</b>			
0.130 ± 0.006	SOKHOYAN	15A	DPWA Multichannel
0.141 ± 0.002	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
0.143 ± 0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
0.147 ± 0.010	AHRENS	02	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.146 ± 0.001	SHKLYAR	13	DPWA Multichannel
0.131 ± 0.010	ANISOVICH	12A	DPWA Multichannel
0.127 ± 0.003	SHRESTHA	12A	DPWA Multichannel
0.161	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
0.151	PENNER	02D	DPWA Multichannel
0.130 ± 0.020 ± 0.015	<sup>2</sup> MUKHOPAD... 98		$\gamma p \rightarrow \eta p$

# Baryon Particle Listings

## $N(1520)$ , $N(1535)$

### $N(1520) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>−0.050±0.010 OUR ESTIMATE</b>			
−0.049±0.008	ANISOVICH	13b	DPWA Multichannel
−0.046±0.006	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.038±0.003	SHRESTHA	12A	DPWA Multichannel
−0.077	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
−0.084	PENNER	02D	DPWA Multichannel

### $N(1520) \rightarrow n\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>−0.115±0.010 OUR ESTIMATE</b>			
−0.113±0.012	ANISOVICH	13b	DPWA Multichannel
−0.115±0.005	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.101±0.004	SHRESTHA	12A	DPWA Multichannel
−0.154	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
−0.159	PENNER	02D	DPWA Multichannel

### $N(1520)$ FOOTNOTES

- <sup>1</sup> Fit to the amplitudes of HOEHLER 79.  
<sup>2</sup> MUKHOPADHYAY 98 uses an effective Lagrangian approach to analyze  $\eta$  photoproduction data. The ratio of the  $A_{3/2}$  and  $A_{1/2}$  amplitudes is determined, with less model dependence than the amplitudes themselves, to be  $A_{3/2}/A_{1/2} = -2.5 \pm 0.5 \pm 0.4$ .

### $N(1520)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982). For very early references, see Reviews of Modern Physics **37** 633 (1965).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
CHEN	12A	PR C86 015206	W. Chen <i>et al.</i>	(DUKE, GWU, MSST, ITEP+)
SHRESTHA	12A	PR C86 065203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
THOMA	08	PL B659 87	U. Thoma <i>et al.</i>	(CB-ELSA Collab.)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER	07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
ARNDT	05	PR C72 045202	R.A. Arndt <i>et al.</i>	(GWU, PNPI)
AHRENS	02	PRL 88 232002	J. Ahrens <i>et al.</i>	(Mainz MAMI GDH/A2 Collab.)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
TIATOR	99	PR C60 035210	L. Tiator <i>et al.</i>	
MUKHOPAD...	98	PL B444 7	N.C. Mukhopadhyay, N. Mathur	
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)UP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)UP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT)UP
Also		Toronto Conf. 3	R. Koch	(KARLT)UP

## $N(1535)$ 1/2<sup>−</sup>

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^-) \text{ Status: } ***$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $N(1535)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1490 to 1530 (<math>\approx 1510</math>) OUR ESTIMATE</b>			
1500 ± 4	SOKHOYAN	15A	DPWA Multichannel
1509 ± 4±2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1502	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1487	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1510±50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1490	SHKLYAR	13	DPWA Multichannel
1501± 4	ANISOVICH	12A	DPWA Multichannel
1515	SHRESTHA	12A	DPWA Multichannel
1521±14	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1525	VRANA	00	DPWA Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>90 to 250 (<math>\approx 170</math>) OUR ESTIMATE</b>			
128 ± 9	SOKHOYAN	15A	DPWA Multichannel
118 ± 9±2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
95	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
260±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

100	SHKLYAR	13	DPWA Multichannel
134±11	ANISOVICH	12A	DPWA Multichannel
123	SHRESTHA	12A	DPWA Multichannel
190±28	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
102	VRANA	00	DPWA Multichannel

### $N(1535)$ ELASTIC POLE RESIDUE

#### MODULUS $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>50±20 OUR ESTIMATE</b>			
29 ± 4	SOKHOYAN	15A	DPWA Multichannel
22 ± 2±0.4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
16	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
120±40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
15	SHKLYAR	13	DPWA Multichannel
31 ± 4	ANISOVICH	12A	DPWA Multichannel
68	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

#### PHASE $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>−15±15 OUR ESTIMATE</b>			
−20±10	SOKHOYAN	15A	DPWA Multichannel
− 5 ± 5±3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−16	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
+15±45	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−51	SHKLYAR	13	DPWA Multichannel
−29 ± 5	ANISOVICH	12A	DPWA Multichannel
12	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

### $N(1535)$ INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow N(1535) \rightarrow N\eta$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
43±3	−76 ± 5	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1535) \rightarrow \Delta\pi, D\text{-wave}$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
11±2	160 ± 20	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
12±3	145 ± 17	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1535) \rightarrow N\sigma$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.16±0.07	25 ± 40	SOKHOYAN	15A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1535) \rightarrow N(1440)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.21±0.14	−45 ± 50	SOKHOYAN	15A	DPWA Multichannel

### $N(1535)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1525 to 1545 (<math>\approx 1535</math>) OUR ESTIMATE</b>			
1517 ± 4	SOKHOYAN	15A	DPWA Multichannel
1526 ± 2	SHKLYAR	13	DPWA Multichannel
1547.0± 0.7	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1550 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1526 ± 7	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1519 ± 5	ANISOVICH	12A	DPWA Multichannel
1538 ± 1	SHRESTHA	12A	DPWA Multichannel
1553 ± 8	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1546.7± 2.2	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$
1526 ± 2	PENNER	02C	DPWA Multichannel
1530 ± 10	BAI	01b	BES $J/\psi \rightarrow p\bar{p}\eta$
1522 ± 11	THOMPSON	01	CLAS $\gamma^*p \rightarrow p\eta$
1542 ± 3	VRANA	00	DPWA Multichannel
1532 ± 5	ARMSTRONG	99b	DPWA $\gamma^*p \rightarrow p\eta$

### $N(1535)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>125 to 175 (<math>\approx 150</math>) OUR ESTIMATE</b>			
120 ± 10	SOKHOYAN	15A	DPWA Multichannel
131 ± 12	SHKLYAR	13	DPWA Multichannel
188.4± 3.8	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
240 ± 80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
120 ± 20	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

See key on page 601

# Baryon Particle Listings

## $N(1535)$ , $N(1650)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

128 ±14	ANISOVICH	12A	DPWA	Multichannel
141 ± 4	SHRESTHA	12A	DPWA	Multichannel
182 ±25	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
129 ± 8	PENNER	02C	DPWA	Multichannel
95 ±25	BAI	01B	BES	$J/\psi \rightarrow p\bar{p}\eta$
143 ±18	THOMPSON	01	CLAS	$\gamma^*p \rightarrow p\eta$
112 ±19	VRANA	00	DPWA	Multichannel
154 ±20	ARMSTRONG	99B	DPWA	$\gamma^*p \rightarrow p\eta$

### $N(1535)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	35–55 %
$\Gamma_2$ $N\eta$	32–52 %
$\Gamma_3$ $N\pi\pi$	3–14 %
$\Gamma_4$ $\Delta(1232)\pi$	
$\Gamma_5$ $\Delta(1232)\pi, D\text{-wave}$	1–4 %
$\Gamma_6$ $N\sigma$	2–10 %
$\Gamma_7$ $N(1440)\pi$	5–12 %
$\Gamma_8$ $p\gamma$ , helicity=1/2	0.15–0.30 %
$\Gamma_9$ $n\gamma$ , helicity=1/2	0.01–0.25 %

### $N(1535)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
52 ± 5	SOKHOYAN	15A	DPWA	Multichannel
35 ± 3	SHKLYAR	13	DPWA	Multichannel
35.5 ± 0.2	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
50 ±10	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
38 ± 4	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
54 ± 5	ANISOVICH	12A	DPWA	Multichannel
37 ± 1	SHRESTHA	12A	DPWA	Multichannel
46 ± 7	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
36 ± 1	PENNER	02C	DPWA	Multichannel
35 ± 8	VRANA	00	DPWA	Multichannel

$\Gamma(N\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
42±10 OUR ESTIMATE				
58 ± 4	SHKLYAR	13	DPWA	Multichannel
33 ± 5	ANISOVICH	12A	DPWA	Multichannel
53 ± 1	PENNER	02C	DPWA	Multichannel
51 ± 5	VRANA	00	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
41 ± 2	SHRESTHA	12A	DPWA	Multichannel
50 ± 7	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$

$\Gamma(N\eta)/\Gamma(N\pi)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma_1$
VALUE				
0.95±0.03	AZNAURYAN	09	CLAS	$\pi, \eta$ electroproduction

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE (%)				
2.5 ±1.5	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.5 ±1.5	ANISOVICH	12A	DPWA	Multichannel
1.8 ±0.8	SHRESTHA	12A	DPWA	Multichannel
1 ±1	VRANA	00	DPWA	Multichannel

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
VALUE (%)				
6 ±4	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.5 ±0.5	SHRESTHA	12A	DPWA	Multichannel
2 ±1	VRANA	00	DPWA	Multichannel

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE (%)				
12 ±8	SOKHOYAN	15A	DPWA	Multichannel
8 ±2	STAROSTIN	03	$\pi^-p \rightarrow n3\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1	SHRESTHA	12A	DPWA	Multichannel
10 ±9	VRANA	00	DPWA	Multichannel

### $N(1535)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $N(1535) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

<u>MODULUS (<math>\text{GeV}^{-1/2}</math>)</u>	<u>PHASE (<math>^\circ</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
0.114 $\pm$ 0.008	10 $\pm$ 5	SOKHOYAN	15A	DPWA	Multichannel

### $N(1535)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $N(1535) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
+0.115 ±0.015 OUR ESTIMATE			
0.101 ±0.007	SOKHOYAN	15A	DPWA Multichannel
0.128 ±0.004	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
0.091 ±0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.091 ±0.004	SHKLYAR	13	DPWA Multichannel
0.105 ±0.010	ANISOVICH	12A	DPWA Multichannel
0.059 ±0.003	SHRESTHA	12A	DPWA Multichannel
0.066	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
0.090	PENNER	02D	DPWA Multichannel

#### $N(1535) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
-0.075 ±0.020 OUR ESTIMATE			
-0.093 ±0.011	ANISOVICH	13B	DPWA Multichannel
-0.058 ±0.006	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.049 ±0.003	SHRESTHA	12A	DPWA Multichannel
-0.051	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
-0.024	PENNER	02D	DPWA Multichannel

#### $N(1535) \rightarrow N\gamma$ , ratio $A_{1/2}^n/A_{1/2}^p$

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN
−0.84 ±0.15	MUKHOPAD... 95B	IPWA

### $N(1535)$ FOOTNOTES

- <sup>1</sup> Fit to the amplitudes of HOEHLER 79.  
<sup>2</sup> This STAROSTIN 03 value is an estimate made using simplest assumptions.

### $N(1535)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN 15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG 14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC 14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH 13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR 13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH 12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
CHEN 12A	PR C86 015206	W. Chen <i>et al.</i>	(DUKE, GWU, MSST, ITEP+)
SHRESTHA 12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN 12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
BATINIC 10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
AZNAURYAN 09	PR C80 055203	I.G. Aznauryan <i>et al.</i>	(JLab CLAS Collab.)
DRECHSEL 07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER 07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT 06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
ARNDT 04	PR C69 035213	R.A. Arndt <i>et al.</i>	(GWU, TRIU)
STAROSTIN 03	PR C67 068201	A. Starostin <i>et al.</i>	(BNL Crystal Ball Collab.)
PENNER 02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER 02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
BAI 01B	PL B510 75	J.Z. Bai <i>et al.</i>	(BES Collab.)
THOMPSON 01	PRL 86 1702	R. Thompson <i>et al.</i>	(JLab CLAS Collab.)
VRANA 00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
ARMSTRONG 99B	PR D60 052004	C.S. Armstrong <i>et al.</i>	
MUKHOPAD... 95B	PL B364 1	N.C. Mukhopadhyay, J.F. Zhang, M. Benmerrouche	
HOEHLER 93	$\pi N$ Newsletter 9 1	G. Hohlner	(KARL)
CUTKOSKY 80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)IJP
Also	PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)IJP
HOEHLER 79	PDAT 12-1	G. Hohlner <i>et al.</i>	(KARL)IJP
Also	Toronto Conf. 3	R. Koch	(KARL)IJP

## $N(1650) 1/2^-$

$I(J^P) = \frac{1}{2}(\frac{1}{2}^-)$  Status: \* \* \* \*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $N(1650)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1640 to 1670 ( $\approx 1655$ ) OUR ESTIMATE			
1652 $\pm$ 7	SOKHOYAN	15A	DPWA Multichannel
1660 $\pm$ 3.5 $\pm$ 1	SVARC	14	L+P $\pi N \rightarrow \pi N$
1648	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1670	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
1640 $\pm$ 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

# Baryon Particle Listings

## N(1650)

• • • We do not use the following data for averages, fits, limits, etc. • • •

1650	SHKLYAR	13	DPWA	Multichannel
1647 ± 6	ANISOVICH	12A	DPWA	Multichannel
1655	SHRESTHA	12A	DPWA	Multichannel
1646 ± 8	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
1663	VRANA	00	DPWA	Multichannel

### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 100 to 170 (≈ 135) OUR ESTIMATE

102 ± 8	SOKHOYAN	15A	DPWA	Multichannel
167 ± 8 ± 2	1 SVARC	14	L+P	$\pi N \rightarrow \pi N$
80	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
163	HOEHLER	93	ARGD	$\pi N \rightarrow \pi N$
150 ± 30	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

89	SHKLYAR	13	DPWA	Multichannel
103 ± 8	ANISOVICH	12A	DPWA	Multichannel
123	SHRESTHA	12A	DPWA	Multichannel
204 ± 17	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
240	VRANA	00	DPWA	Multichannel

### N(1650) ELASTIC POLE RESIDUE

#### MODULUS |r|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 20 to 50 (≈ 35) OUR ESTIMATE

27 ± 6	SOKHOYAN	15A	DPWA	Multichannel
47 ± 3 ± 1	1 SVARC	14	L+P	$\pi N \rightarrow \pi N$
14	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
39	HOEHLER	93	ARGD	$\pi N \rightarrow \pi N$
60 ± 10	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

19	SHKLYAR	13	DPWA	Multichannel
24 ± 3	ANISOVICH	12A	DPWA	Multichannel
100	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$

#### PHASE θ

VALUE (°)	DOCUMENT ID	TECN	COMMENT
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#### 50 to 80 (≈ 70) OUR ESTIMATE

−60 ± 20	SOKHOYAN	15A	DPWA	Multichannel
−47 ± 3 ± 1	1 SVARC	14	L+P	$\pi N \rightarrow \pi N$
−69	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
−37	HOEHLER	93	ARGD	$\pi N \rightarrow \pi N$
−75 ± 25	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

−46	SHKLYAR	13	DPWA	Multichannel
−75 ± 12	ANISOVICH	12A	DPWA	Multichannel
−65	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$

### N(1650) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow N(1650) \rightarrow N\eta$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.29 ± 0.03	134 ± 10	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1650) \rightarrow \Lambda K$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.23 ± 0.09	85 ± 9	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1650) \rightarrow \Delta\pi, D\text{-wave}$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.19 ± 0.06	−30 ± 20	SOKHOYAN	15A	DPWA Multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.23 ± 0.04	−30 ± 20	ANISOVICH	12A	DPWA Multichannel
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#### Normalized residue in $N\pi \rightarrow N(1650) \rightarrow N\sigma$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.20 ± 0.15	undefined	SOKHOYAN	15A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1650) \rightarrow N(1440)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.30 ± 0.17	undefined	SOKHOYAN	15A	DPWA Multichannel

### N(1650) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 1645 to 1670 (≈ 1655) OUR ESTIMATE

1654 ± 6	SOKHOYAN	15A	DPWA	Multichannel
1665 ± 2	SHKLYAR	13	DPWA	Multichannel
1634.7 ± 1.1	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
1650 ± 30	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
1670 ± 8	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1651 ± 6	ANISOVICH	12A	DPWA	Multichannel
1664 ± 2	SHRESTHA	12A	DPWA	Multichannel
1652 ± 9	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
1665 ± 2	PENNER	02c	DPWA	Multichannel
1647 ± 20	BAI	01B	BES	$J/\psi \rightarrow p\bar{p}\eta$
1689 ± 12	VRANA	00	DPWA	Multichannel

### N(1650) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 110 to 170 (≈ 140) OUR ESTIMATE

102 ± 8	SOKHOYAN	15A	DPWA	Multichannel
147 ± 14	SHKLYAR	13	DPWA	Multichannel
115.4 ± 2.8	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
150 ± 40	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
180 ± 20	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

104 ± 10	ANISOVICH	12A	DPWA	Multichannel
126 ± 3	SHRESTHA	12A	DPWA	Multichannel
202 ± 16	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
138 ± 7	PENNER	02c	DPWA	Multichannel
145 +80 −45	BAI	01B	BES	$J/\psi \rightarrow p\bar{p}\eta$
202 ± 40	VRANA	00	DPWA	Multichannel

### N(1650) DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	50–70 %
$\Gamma_2$ $N\eta$	14–22 %
$\Gamma_3$ $\Lambda K$	5–15 %
$\Gamma_4$ $N\pi\pi$	8–36 %
$\Gamma_5$ $\Delta(1232)\pi$	
$\Gamma_6$ $\Delta(1232)\pi, D\text{-wave}$	6–18 %
$\Gamma_7$ $N\sigma$	2–18 %
$\Gamma_8$ $N(1440)\pi$	6–26 %
$\Gamma_9$ $p\gamma, \text{ helicity}=1/2$	0.04–0.20 %
$\Gamma_{10}$ $n\gamma, \text{ helicity}=1/2$	0.003–0.17 %

### N(1650) BRANCHING RATIOS

#### $\Gamma(N\pi)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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#### 50 to 70 (≈ 60) OUR ESTIMATE

51 ± 4	SOKHOYAN	15A	DPWA	Multichannel
74 ± 3	SHKLYAR	13	DPWA	Multichannel
65 ± 10	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
61 ± 4	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

51 ± 4	ANISOVICH	12A	DPWA	Multichannel
57 ± 2	SHRESTHA	12A	DPWA	Multichannel
79 ± 6	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
100	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
65 ± 4	PENNER	02c	DPWA	Multichannel
74 ± 2	VRANA	00	DPWA	Multichannel

#### $\Gamma(N\eta)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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1 ± 2	SHKLYAR	13	DPWA	Multichannel
18 ± 4	ANISOVICH	12A	DPWA	Multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •

21 ± 2	SHRESTHA	12A	DPWA	Multichannel
13 ± 5	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
1.0 ± 0.6	PENNER	02c	DPWA	Multichannel
6 ± 1	VRANA	00	DPWA	Multichannel

#### $\Gamma(\Lambda K)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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#### 5 to 15 OUR ESTIMATE

10 ± 5	ANISOVICH	12A	DPWA	Multichannel
4 ± 1	SHKLYAR	05	DPWA	Multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •

8 ± 1	SHRESTHA	12A	DPWA	Multichannel
2.7 ± 0.4	PENNER	02c	DPWA	Multichannel



See key on page 601

Baryon Particle Listings  
*N*(1650), *N*(1675)

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
12±6	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
19±9	ANISOVICH	12A	DPWA Multichannel
7±2	SHRESTHA	12A	DPWA Multichannel
2±1	VRANA	00	DPWA Multichannel

$\Gamma(N\sigma)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
10±8	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 1	SHRESTHA	12A	DPWA Multichannel
1±1	VRANA	00	DPWA Multichannel

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
16±10	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 1	SHRESTHA	12A	DPWA Multichannel
3± 1	VRANA	00	DPWA Multichannel

*N*(1650) PHOTON DECAY AMPLITUDES AT THE POLE

*N*(1650) → *p*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.032±0.006	−2 ± 11	SOKHOYAN	15A	DPWA Multichannel

*N*(1650) BREIT-WIGNER PHOTON DECAY AMPLITUDES

*N*(1650) → *p*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>+0.045±0.010 OUR ESTIMATE</b>			
0.032±0.006	SOKHOYAN	15A	DPWA Multichannel
0.055±0.030	WORKMAN	12A	DPWA $\gamma N \rightarrow \pi N$
0.022±0.007	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.063±0.006	SHKLYAR	13	DPWA Multichannel
0.033±0.007	ANISOVICH	12A	DPWA Multichannel
0.030±0.003	SHRESTHA	12A	DPWA Multichannel
0.033	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
0.049	PENNER	02D	DPWA Multichannel

*N*(1650) → *n*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>−0.050±0.020 OUR ESTIMATE</b>			
0.025±0.020	ANISOVICH	13B	DPWA Multichannel
−0.040±0.010	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.011±0.002	SHRESTHA	12A	DPWA Multichannel
0.009	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
−0.011	PENNER	02D	DPWA Multichannel

*N*(1650) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

*N*(1650) REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
CHEN	12A	PR C86 015206	W. Chen <i>et al.</i>	(DUKE, GWU, MSST, ITEP+)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER	07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	
SHKLYAR	05	PR C72 015210	V. Shklyar, H. Lenske, U. Mosel	(GIES)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
BAI	01B	PL B510 75	J.Z. Bai <i>et al.</i>	(BES Collab.)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	93	π N Newsletter 9 1	G. Hohlner	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohlner <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

*N*(1675) 5/2<sup>−</sup>

$I(J^P) = \frac{1}{2}(\frac{5}{2}^-)$  Status: \* \* \* \*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

*N*(1675) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1655 to 1665 (≈ 1660) OUR ESTIMATE</b>			
1655 ± 4	SOKHOYAN	15A	DPWA Multichannel
1654 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1657	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1656	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
1660±10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1640	SHKLYAR	13	DPWA Multichannel
1654 ± 4	ANISOVICH	12A	DPWA Multichannel
1656	SHRESTHA	12A	DPWA Multichannel
1658 ± 9	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1674	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>125 to 150 (≈ 135) OUR ESTIMATE</b>			
147 ± 5	SOKHOYAN	15A	DPWA Multichannel
125 ± 3±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
139	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
126	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
140±10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
108	SHKLYAR	13	DPWA Multichannel
151 ± 5	ANISOVICH	12A	DPWA Multichannel
128	SHRESTHA	12A	DPWA Multichannel
137 ± 7	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
120	VRANA	00	DPWA Multichannel

*N*(1675) ELASTIC POLE RESIDUE

MODULUS |*r*|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>27±5 OUR ESTIMATE</b>			
28±1	SOKHOYAN	15A	DPWA Multichannel
23±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
27	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
23	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
31±5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
20	SHKLYAR	13	DPWA Multichannel
28±1	ANISOVICH	12A	DPWA Multichannel
25	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

PHASE θ

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>−25 ± 6 OUR ESTIMATE</b>			
−24 ± 4	SOKHOYAN	15A	DPWA Multichannel
−25 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−21	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−22	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
−30±10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−49	SHKLYAR	13	DPWA Multichannel
−26 ± 4	ANISOVICH	12A	DPWA Multichannel
−16	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

*N*(1675) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{\text{pole}}/2$ .

Normalized residue in  $N\pi \rightarrow N(1675) \rightarrow \Delta\pi, D\text{-wave}$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
33±4	90 ± 15	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
33±5	82 ± 10	ANISOVICH	12A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow N(1675) \rightarrow N\sigma$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
13±3	125 ± 20	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
15±4	132 ± 18	ANISOVICH	12A	DPWA Multichannel

Baryon Particle Listings

*N*(1675)

*N*(1675) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1670 to 1680 (<math>\approx 1675</math>) OUR ESTIMATE</b>			
1663 $\pm$ 4	SOKHOYAN	15A	DPWA Multichannel
1666 $\pm$ 2	SHKLYAR	13	DPWA Multichannel
1674.1 $\pm$ 0.2	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1675 $\pm$ 10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1679 $\pm$ 8	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1664 $\pm$ 5	ANISOVICH	12A	DPWA Multichannel
1679 $\pm$ 1	SHRESTHA	12A	DPWA Multichannel
1679 $\pm$ 9	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1685 $\pm$ 4	VRANA	00	DPWA Multichannel

*N*(1675) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>130 to 165 (<math>\approx 150</math>) OUR ESTIMATE</b>			
146 $\pm$ 6	SOKHOYAN	15A	DPWA Multichannel
148 $\pm$ 1	SHKLYAR	13	DPWA Multichannel
146.5 $\pm$ 1.0	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
160 $\pm$ 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
120 $\pm$ 15	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
152 $\pm$ 7	ANISOVICH	12A	DPWA Multichannel
145 $\pm$ 4	SHRESTHA	12A	DPWA Multichannel
152 $\pm$ 8	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
131 $\pm$ 10	VRANA	00	DPWA Multichannel

*N*(1675) DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	35–45 %
$\Gamma_2$ $N\eta$	< 1 %
$\Gamma_3$ $N\pi\pi$	25–45 %
$\Gamma_4$ $\Delta(1232)\pi$	
$\Gamma_5$ $\Delta(1232)\pi, D\text{-wave}$	23–37 %
$\Gamma_6$ $N\sigma$	3–7 %
$\Gamma_7$ $p\gamma$	0–0.02 %
$\Gamma_8$ $p\gamma$ , helicity=1/2	0–0.01 %
$\Gamma_9$ $p\gamma$ , helicity=3/2	0–0.01 %
$\Gamma_{10}$ $n\gamma$	0–0.15 %
$\Gamma_{11}$ $n\gamma$ , helicity=1/2	0–0.05 %
$\Gamma_{12}$ $n\gamma$ , helicity=3/2	0–0.10 %

*N*(1675) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>35 to 45 OUR ESTIMATE</b>				
41 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel	
41 $\pm$ 1	SHKLYAR	13	DPWA Multichannel	
39.3 $\pm$ 0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$	
38 $\pm$ 5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
38 $\pm$ 3	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
40 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel	
38.6 $\pm$ 0.6	SHRESTHA	12A	DPWA Multichannel	
35 $\pm$ 4	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
35 $\pm$ 1	VRANA	00	DPWA Multichannel	

$\Gamma(N\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
<b>VALUE (%)</b>				
0 $\pm$ 1	SHKLYAR	13	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1	SHRESTHA	12A	DPWA Multichannel	
0.1 $\pm$ 0.1	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
3 $\pm$ 3	THOMA	08	DPWA Multichannel	
0 $\pm$ 1	VRANA	00	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
<b>VALUE (%)</b>				
30 $\pm$ 7	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
33 $\pm$ 8	ANISOVICH	12A	DPWA Multichannel	
46 $\pm$ 1	SHRESTHA	12A	DPWA Multichannel	
63 $\pm$ 2	VRANA	00	DPWA Multichannel	

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
<b>VALUE (%)</b>				
5 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel	

*N*(1675) PHOTON DECAY AMPLITUDES AT THE POLE

*N*(1675)  $\rightarrow p\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

MODULUS (GeV $^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
0.022 $\pm$ 0.003	–12 $\pm$ 7	SOKHOYAN	15A	DPWA Multichannel

*N*(1675)  $\rightarrow p\gamma$ , helicity-3/2 amplitude  $A_{3/2}$

MODULUS (GeV $^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
0.028 $\pm$ 0.006	–17 $\pm$ 6	SOKHOYAN	15A	DPWA Multichannel

*N*(1675) BREIT-WIGNER PHOTON DECAY AMPLITUDES

*N*(1675)  $\rightarrow p\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>+0.019<math>\pm</math>0.008 OUR ESTIMATE</b>			
0.022 $\pm$ 0.003	SOKHOYAN	15A	DPWA Multichannel
0.013 $\pm$ 0.001	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
0.018 $\pm$ 0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.009 $\pm$ 0.001	SHKLYAR	13	DPWA Multichannel
0.024 $\pm$ 0.003	ANISOVICH	12A	DPWA Multichannel
0.011 $\pm$ 0.001	SHRESTHA	12A	DPWA Multichannel
0.015	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1675)  $\rightarrow p\gamma$ , helicity-3/2 amplitude  $A_{3/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>+0.020<math>\pm</math>0.005 OUR ESTIMATE</b>			
0.027 $\pm$ 0.006	SOKHOYAN	15A	DPWA Multichannel
0.016 $\pm$ 0.001	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
0.021 $\pm$ 0.001	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.021 $\pm$ 0.001	SHKLYAR	13	DPWA Multichannel
0.025 $\pm$ 0.007	ANISOVICH	12A	DPWA Multichannel
0.020 $\pm$ 0.001	SHRESTHA	12A	DPWA Multichannel
0.022	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1675)  $\rightarrow n\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>–0.060<math>\pm</math>0.005 OUR ESTIMATE</b>			
–0.060 $\pm$ 0.007	ANISOVICH	13B	DPWA Multichannel
–0.058 $\pm$ 0.002	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–0.040 $\pm$ 0.004	SHRESTHA	12A	DPWA Multichannel
–0.062	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1675)  $\rightarrow n\gamma$ , helicity-3/2 amplitude  $A_{3/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>–0.085<math>\pm</math>0.010 OUR ESTIMATE</b>			
–0.088 $\pm$ 0.010	ANISOVICH	13B	DPWA Multichannel
–0.080 $\pm$ 0.005	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–0.068 $\pm$ 0.004	SHRESTHA	12A	DPWA Multichannel
–0.084	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1675) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

*N*(1675) REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	(GIES)
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(BONN, PNPI)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(DUKE, GWU, MSST, ITEP+)
CHEN	12A	PR C86 015206	W. Chen <i>et al.</i>	(KSU)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(ZAGR)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(CB-ELSA Collab.)
THOMA	08	PL B659 87	U. Thoma <i>et al.</i>	(MAINZ, JINR)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(JLab CLAS Collab.)
DUGGER	07	PR C76 025211	M. Dugger <i>et al.</i>	(GWU)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(PITT, ANL)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(KARL)
HOEHLER	79	$\pi N$ Newsletter 9 1	G. H��hler	(CMU, LBL) IUP
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(KARLT) IUP
HOEHLER	79	PDAT 12-1	G. H��hler <i>et al.</i>	(KARLT) IUP
Also		Toronto Conf. 3	R. Koch	



Baryon Particle Listings

*N*(1680), *N*(1700)

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$	
VALUE (%)	DOCUMENT ID	TECN	COMMENT		
7 ± 3	SOKHOYAN	15A	DPWA Multichannel		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
5 ± 3	ANISOVICH	12A	DPWA Multichannel		
10.5 ± 0.9	SHRESTHA	12A	DPWA Multichannel		
14 ± 3	VRANA	00	DPWA Multichannel		

$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$	
VALUE (%)	DOCUMENT ID	TECN	COMMENT		
10 ± 3	SOKHOYAN	15A	DPWA Multichannel		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
10 ± 3	ANISOVICH	12A	DPWA Multichannel		
1.0 ± 0.1	SHRESTHA	12A	DPWA Multichannel		
1 ± 1	VRANA	00	DPWA Multichannel		

$\Gamma(N\sigma)/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$	
VALUE (%)	DOCUMENT ID	TECN	COMMENT		
14 ± 5	SOKHOYAN	15A	DPWA Multichannel		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
14 ± 7	ANISOVICH	12A	DPWA Multichannel		
9.4 ± 0.8	SHRESTHA	12A	DPWA Multichannel		
9 ± 1	VRANA	00	DPWA Multichannel		

*N*(1680) PHOTON DECAY AMPLITUDES AT THE POLE

*N*(1680) → *p* $\gamma$ , helicity-1/2 amplitude *A*<sub>1/2</sub>

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
−0.013 ± 0.003	−20 ± 17	SOKHOYAN	15A	DPWA Multichannel

*N*(1680) → *p* $\gamma$ , helicity-3/2 amplitude *A*<sub>3/2</sub>

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.135 ± 0.005	1 ± 3	SOKHOYAN	15A	DPWA Multichannel

*N*(1680) BREIT-WIGNER PHOTON DECAY AMPLITUDES

*N*(1680) → *p* $\gamma$ , helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.015 ± 0.006 OUR ESTIMATE			
−0.015 ± 0.002	SOKHOYAN	15A	DPWA Multichannel
−0.007 ± 0.002	WORKMAN	12A	DPWA $\gamma N \rightarrow \pi N$
−0.017 ± 0.001	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.003 ± 0.001	SHKLYAR	13	DPWA Multichannel
−0.013 ± 0.003	ANISOVICH	12A	DPWA Multichannel
−0.017 ± 0.001	SHRESTHA	12A	DPWA Multichannel
−0.025	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1680) → *p* $\gamma$ , helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
+0.133 ± 0.012 OUR ESTIMATE			
0.136 ± 0.005	SOKHOYAN	15A	DPWA Multichannel
0.140 ± 0.002	WORKMAN	12A	DPWA $\gamma N \rightarrow \pi N$
0.134 ± 0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.116 ± 0.001	SHKLYAR	13	DPWA Multichannel
0.135 ± 0.006	ANISOVICH	12A	DPWA Multichannel
0.136 ± 0.001	SHRESTHA	12A	DPWA Multichannel
0.134	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1680) → *n* $\gamma$ , helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
+0.029 ± 0.010 OUR ESTIMATE			
0.034 ± 0.006	ANISOVICH	13B	DPWA Multichannel
0.026 ± 0.004	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.029 ± 0.002	SHRESTHA	12A	DPWA Multichannel
0.028	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1680) → *n* $\gamma$ , helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.033 ± 0.009 OUR ESTIMATE			
−0.044 ± 0.009	ANISOVICH	13B	DPWA Multichannel
−0.029 ± 0.002	CHEN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.059 ± 0.002	SHRESTHA	12A	DPWA Multichannel
−0.038	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

*N*(1680) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

*N*(1680) REFERENCES

For early references, see Physics Letters **111B** 1 (1982). For very early references, see Reviews of Modern Physics **37** 633 (1965).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
CHEN	12A	PR C86 015206	W. Chen <i>et al.</i>	(DUKE, GWU, MSST, ITEP+)
SHRESTHA	12A	PR C86 065203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
THOMA	08	PL B659 87	U. Thoma <i>et al.</i>	(CB-ELSA Collab.)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER	07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
TIATOR	99	PR C60 035210	L. Tiator <i>et al.</i>	
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

*N*(1700) 3/2<sup>−</sup>

$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$  Status: \*\*\*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

*N*(1700) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1650 to 1750 (≈ 1700) OUR ESTIMATE			
1780 ± 35	SOKHOYAN	15A	DPWA Multichannel
1757 ± 4 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1700	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1660 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1770 ± 40	ANISOVICH	12A	DPWA Multichannel
1662	SHRESTHA	12A	DPWA Multichannel
1806 ± 23	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1704	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
100 to 300 OUR ESTIMATE			
420 ± 140	SOKHOYAN	15A	DPWA Multichannel
136 ± 7 ± 4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
120	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
90 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
420 ± 180	ANISOVICH	12A	DPWA Multichannel
55	SHRESTHA	12A	DPWA Multichannel
129 ± 33	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
156	VRANA	00	DPWA Multichannel

*N*(1700) ELASTIC POLE RESIDUE

MODULUS |*r*|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
5 to 50 OUR ESTIMATE			
60 ± 30	SOKHOYAN	15A	DPWA Multichannel
7 ± 1 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
5	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
6 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
50 ± 40	ANISOVICH	12A	DPWA Multichannel
7	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−120 to 20 OUR ESTIMATE			
−115 ± 30	SOKHOYAN	15A	DPWA Multichannel
−113 ± 4 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
0 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−100 ± 40	ANISOVICH	12A	DPWA Multichannel
− 34	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

See key on page 601

# Baryon Particle Listings

## $N(1700)$

### $N(1700)$ INELASTIC POLE RESIDUE

The "normalized residue" is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow N(1700) \rightarrow \Delta\pi$ , S-wave

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$0.33 \pm 0.10$	$-70 \pm 25$	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.34 \pm 0.21$	$-60 \pm 40$	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1700) \rightarrow \Delta\pi$ , D-wave

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$0.10 \pm 0.06$	$75 \pm 30$	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.08 \pm 0.06$	$90 \pm 35$	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1700) \rightarrow N\sigma$

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$0.13 \pm 0.08$	$-100 \pm 35$	SOKHOYAN	15A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1700) \rightarrow N(1440)\pi$

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$0.13 \pm 0.05$	$40 \pm 35$	SOKHOYAN	15A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1700) \rightarrow N(1520)\pi$ , P-wave

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$0.07 \pm 0.03$	$160 \pm 45$	SOKHOYAN	15A	DPWA Multichannel

### $N(1700)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1650 to 1750 (<math>\approx 1700</math>) OUR ESTIMATE</b>			
$1800 \pm 35$	SOKHOYAN	15A	DPWA Multichannel
$1675 \pm 25$	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
$1731 \pm 15$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1790 \pm 40$	ANISOVICH	12A	DPWA Multichannel
$1665 \pm 3$	SHRESTHA	12A	DPWA Multichannel
$1817 \pm 22$	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
$1736 \pm 33$	VRANA	00	DPWA Multichannel

### $N(1700)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>100 to 250 (<math>\approx 150</math>) OUR ESTIMATE</b>			
$400 \pm 100$	SOKHOYAN	15A	DPWA Multichannel
$90 \pm 40$	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
$110 \pm 30$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$390 \pm 140$	ANISOVICH	12A	DPWA Multichannel
$56 \pm 8$	SHRESTHA	12A	DPWA Multichannel
$134 \pm 37$	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
$175 \pm 133$	VRANA	00	DPWA Multichannel

### $N(1700)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	7–17 %
$\Gamma_2$ $N\eta$	seen
$\Gamma_3$ $N\pi\pi$	60–90 %
$\Gamma_4$ $\Delta(1232)\pi$	55–85 %
$\Gamma_5$ $\Delta(1232)\pi$ , S-wave	50–80 %
$\Gamma_6$ $\Delta(1232)\pi$ , D-wave	4–14 %
$\Gamma_7$ $N(1440)\pi$	3–11 %
$\Gamma_8$ $N(1520)\pi$	<4 %
$\Gamma_9$ $N\rho$ , $S=3/2$ , S-wave	seen
$\Gamma_{10}$ $N\sigma$	2–14 %
$\Gamma_{11}$ $p\gamma$	0.01–0.05 %
$\Gamma_{12}$ $p\gamma$ , helicity=1/2	0.0–0.024 %
$\Gamma_{13}$ $p\gamma$ , helicity=3/2	0.002–0.026 %
$\Gamma_{14}$ $n\gamma$	0.01–0.13 %
$\Gamma_{15}$ $n\gamma$ , helicity=1/2	0.0–0.09 %
$\Gamma_{16}$ $n\gamma$ , helicity=3/2	0.01–0.05 %

### $N(1700)$ BRANCHING RATIOS

#### $\Gamma(N\pi)/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>12 <math>\pm</math> 5 OUR ESTIMATE</b>			
$15 \pm 6$	SOKHOYAN	15A	DPWA Multichannel
$11 \pm 5$	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
$8 \pm 3$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$12 \pm 5$	ANISOVICH	12A	DPWA Multichannel
$2.8 \pm 0.5$	SHRESTHA	12A	DPWA Multichannel
$9 \pm 6$	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
$4 \pm 2$	VRANA	00	DPWA Multichannel

#### $\Gamma_1/\Gamma$

#### $\Gamma(N\eta)/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$14 \pm 5$	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
$10 \pm 5$	THOMA	08	DPWA Multichannel
$0 \pm 1$	VRANA	00	DPWA Multichannel

#### $\Gamma_2/\Gamma$

#### $\Gamma(\Delta(1232)\pi, S\text{-wave})/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$65 \pm 15$	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$72 \pm 23$	ANISOVICH	12A	DPWA Multichannel
$31 \pm 9$	SHRESTHA	12A	DPWA Multichannel
$11 \pm 1$	VRANA	00	DPWA Multichannel

#### $\Gamma_5/\Gamma$

#### $\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$9 \pm 5$	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<10	ANISOVICH	12A	DPWA Multichannel
$3 \pm 2$	SHRESTHA	12A	DPWA Multichannel
$79 \pm 56$	VRANA	00	DPWA Multichannel

#### $\Gamma_6/\Gamma$

#### $\Gamma(N(1440)\pi)/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$7 \pm 4$	SOKHOYAN	15A	DPWA Multichannel

#### $\Gamma_7/\Gamma$

#### $\Gamma(N(1520)\pi)/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<4	SOKHOYAN	15A	DPWA Multichannel

#### $\Gamma_8/\Gamma$

#### $\Gamma(N\rho, S=3/2, S\text{-wave})/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$38 \pm 6$	SHRESTHA	12A	DPWA Multichannel
$7 \pm 1$	VRANA	00	DPWA Multichannel

#### $\Gamma_9/\Gamma$

#### $\Gamma(N\sigma)/\Gamma_{total}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
$8 \pm 6$	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$24 \pm 6$	SHRESTHA	12A	DPWA Multichannel
$18 \pm 12$	THOMA	08	DPWA Multichannel
$0 \pm 1$	VRANA	00	DPWA Multichannel

#### $\Gamma_{10}/\Gamma$

### $N(1700)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $N(1700) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$0.047 \pm 0.016$	$75 \pm 30$	SOKHOYAN	15A	DPWA Multichannel

#### $N(1700) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$

MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
$-0.041 \pm 0.014$	$0 \pm 20$	SOKHOYAN	15A	DPWA Multichannel

### $N(1700)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $N(1700) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
$0.041 \pm 0.017$	ANISOVICH	12A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.021 \pm 0.005$	SHRESTHA	12A	DPWA Multichannel

# Baryon Particle Listings

## N(1700), N(1710)

### N(1700) → pγ, helicity-3/2 amplitude A3/2

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.037 ± 0.014	SOKHOYAN 15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.034 ± 0.013	ANISOVICH 12A	DPWA	Multichannel
0.050 ± 0.009	SHRESTHA 12A	DPWA	Multichannel

### N(1700) → nγ, helicity-1/2 amplitude A1/2

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.025 ± 0.010	ANISOVICH 13B	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.049 ± 0.008	SHRESTHA 12A	DPWA	Multichannel

### N(1700) → nγ, helicity-3/2 amplitude A3/2

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.032 ± 0.018	ANISOVICH 13B	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.092 ± 0.014	SHRESTHA 12A	DPWA	Multichannel

### N(1700) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### N(1700) REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN 15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG 14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC 14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH 13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH 12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA 12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
BATINIC 10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
THOMA 08	PL B659 87	U. Thoma <i>et al.</i>	(CB-ELSA Collab.)
VRANA 00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER 93	π N Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY 80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also	PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER 79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also	Toronto Conf. 3	R. Koch	(KARLT) IJP

N(1710) 1/2<sup>+</sup>

I(J<sup>P</sup>) =  $\frac{1}{2}(\frac{1}{2}^+)$  Status: \* \* \* \*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### N(1710) POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1670 to 1770 (≈ 1720) OUR ESTIMATE</b>			
1690 ± 15	SOKHOYAN 15A	DPWA	Multichannel
1770 ± 5 ± 2	<sup>1</sup> SVARC 14	L+P	π N → π N
1690	HOEHLER 93	SPED	π N → π N
1698	CUTKOSKY 90	IPWA	π N → π N
1690 ± 20	CUTKOSKY 80	IPWA	π N → π N
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1690 ± 15	GUTZ 14	DPWA	Multichannel
1670	SHKLYAR 13	DPWA	Multichannel
1687 ± 17	ANISOVICH 12A	DPWA	Multichannel
1644	SHRESTHA 12A	DPWA	Multichannel
1711 ± 15	<sup>2</sup> BATINIC 10	DPWA	π N → Nπ, Nη
1679	VRANA 00	DPWA	Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>80 to 380 (≈ 230) OUR ESTIMATE</b>			
170 ± 20	SOKHOYAN 15A	DPWA	Multichannel
98 ± 8 ± 5	<sup>1</sup> SVARC 14	L+P	π N → π N
200	HOEHLER 93	SPED	π N → π N
88	CUTKOSKY 90	IPWA	π N → π N
80 ± 20	CUTKOSKY 80	IPWA	π N → π N
• • • We do not use the following data for averages, fits, limits, etc. • • •			
170 ± 20	GUTZ 14	DPWA	Multichannel
159	SHKLYAR 13	DPWA	Multichannel
200 ± 25	ANISOVICH 12A	DPWA	Multichannel
104	SHRESTHA 12A	DPWA	Multichannel
174 ± 16	<sup>2</sup> BATINIC 10	DPWA	π N → Nπ, Nη
132	VRANA 00	DPWA	Multichannel

### N(1710) ELASTIC POLE RESIDUE

#### MODULUS |r|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5 to 15 (≈ 8) OUR ESTIMATE</b>			
6 ± 3	SOKHOYAN 15A	DPWA	Multichannel
5 ± 1 ± 1	<sup>1</sup> SVARC 14	L+P	π N → π N
15	HOEHLER 93	SPED	π N → π N
9	CUTKOSKY 90	IPWA	π N → π N
8 ± 2	CUTKOSKY 80	IPWA	π N → π N
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6 ± 3	GUTZ 14	DPWA	Multichannel
11	SHKLYAR 13	DPWA	Multichannel
6 ± 4	ANISOVICH 12A	DPWA	Multichannel
24	<sup>2</sup> BATINIC 10	DPWA	π N → Nπ, Nη

#### PHASE θ

VALUE (°)	DOCUMENT ID	TECN	COMMENT
130 ± 35	SOKHOYAN 15A	DPWA	Multichannel
−104 ± 7 ± 3	<sup>1</sup> SVARC 14	L+P	π N → π N
−167	CUTKOSKY 90	IPWA	π N → π N
175 ± 35	CUTKOSKY 80	IPWA	π N → π N
• • • We do not use the following data for averages, fits, limits, etc. • • •			
120 ± 45	GUTZ 14	DPWA	Multichannel
9	SHKLYAR 13	DPWA	Multichannel
120 ± 70	ANISOVICH 12A	DPWA	Multichannel
20	<sup>2</sup> BATINIC 10	DPWA	π N → Nπ, Nη

### N(1710) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by Γ<sub>pole</sub>/2.

#### Normalized residue in Nπ → N(1710) → Nη

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
12 ± 4	0 ± 45	ANISOVICH 12A	DPWA	Multichannel

#### Normalized residue in Nπ → N(1710) → ΛK

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
17 ± 6	−110 ± 20	ANISOVICH 12A	DPWA	Multichannel

#### Normalized residue in Nπ → N(1710) → N(1535)π

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
10 ± 4	140 ± 40	GUTZ 14	DPWA	Multichannel

### N(1710) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1680 to 1740 (≈ 1710) OUR ESTIMATE</b>			
1715 ± 20	SOKHOYAN 15A	DPWA	Multichannel
1737 ± 17	SHKLYAR 13	DPWA	Multichannel
1700 ± 50	CUTKOSKY 80	IPWA	π N → π N
1723 ± 9	HOEHLER 79	IPWA	π N → π N
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1715 ± 20	GUTZ 14	DPWA	Multichannel
1710 ± 20	ANISOVICH 12A	DPWA	Multichannel
1662 ± 7	SHRESTHA 12A	DPWA	Multichannel
1729 ± 16	<sup>2</sup> BATINIC 10	DPWA	π N → Nπ, Nη
1752 ± 3	PENNER 02C	DPWA	Multichannel
1699 ± 65	VRANA 00	DPWA	Multichannel

### N(1710) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>50 to 250 (≈ 100) OUR ESTIMATE</b>			
175 ± 15	SOKHOYAN 15A	DPWA	Multichannel
368 ± 120	SHKLYAR 13	DPWA	Multichannel
93 ± 30	CUTKOSKY 90	IPWA	π N → π N
90 ± 30	CUTKOSKY 80	IPWA	π N → π N
120 ± 15	HOEHLER 79	IPWA	π N → π N
• • • We do not use the following data for averages, fits, limits, etc. • • •			
175 ± 15	GUTZ 14	DPWA	Multichannel
200 ± 18	ANISOVICH 12A	DPWA	Multichannel
116 ± 17	SHRESTHA 12A	DPWA	Multichannel
180 ± 17	<sup>2</sup> BATINIC 10	DPWA	π N → Nπ, Nη
386 ± 59	PENNER 02C	DPWA	Multichannel
143 ± 100	VRANA 00	DPWA	Multichannel

See key on page 601

Baryon Particle Listings  
*N*(1710), *N*(1720)

*N*(1710) DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	5–20 %
$\Gamma_2$ $N\eta$	10–50 %
$\Gamma_3$ $N\omega$	1–5 %
$\Gamma_4$ $\Lambda K$	5–25 %
$\Gamma_5$ $\Sigma K$	seen
$\Gamma_6$ $N\pi\pi$	seen
$\Gamma_7$ $\Delta(1232)\pi$	
$\Gamma_8$ $\Delta(1232)\pi$ , <i>P</i> -wave	seen
$\Gamma_9$ $N(1535)\pi$	9–21 %
$\Gamma_{10}$ $N\rho$	
$\Gamma_{11}$ $N\rho$ , <i>S</i> =1/2, <i>P</i> -wave	seen
$\Gamma_{12}$ $\rho\gamma$ , helicity=1/2	0.002–0.08 %
$\Gamma_{13}$ $n\gamma$ , helicity=1/2	0.0–0.02%

*N*(1710) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
5 ± 3	SOKHOYAN	15A	DPWA Multichannel	
2 ± 2	SHKLYAR	13	PWA Multichannel	
20 ± 4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
12 ± 4	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5 ± 3	GUTZ	14	DPWA Multichannel	
5 ± 4	ANISOVICH	12A	DPWA Multichannel	
15 ± 4	SHRESTHA	12A	DPWA Multichannel	
22 ± 24	BATINIC	10	DPWA $\pi N \rightarrow N\pi$ , $N\eta$	
14 ± 8	PENNER	02C	DPWA Multichannel	
27 ± 13	VRANA	00	DPWA Multichannel	

$\Gamma(N\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
45 ± 4	SHKLYAR	13	DPWA Multichannel	
17 ± 10	ANISOVICH	12A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
11 ± 7	SHRESTHA	12A	DPWA Multichannel	
6 ± 8	BATINIC	10	DPWA $\pi N \rightarrow N\pi$ , $N\eta$	
36 ± 11	PENNER	02C	DPWA Multichannel	
6 ± 1	VRANA	00	DPWA Multichannel	

$\Gamma(N\omega)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
VALUE (%)				
3 ± 2	SHKLYAR	13	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13 ± 2	PENNER	02C	DPWA Multichannel	

$\Gamma(\Lambda K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE (%)				
5 to 25 OUR ESTIMATE				
23 ± 7	ANISOVICH	12A	DPWA Multichannel	
5 ± 3	SHKLYAR	05	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8 ± 4	SHRESTHA	12A	DPWA Multichannel	
5 ± 2	PENNER	02C	DPWA Multichannel	
10 ± 10	VRANA	00	DPWA Multichannel	

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7 ± 7	PENNER	02C	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6 ± 3	SHRESTHA	12A	DPWA Multichannel	
39 ± 8	VRANA	00	DPWA Multichannel	

$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
VALUE (%)				
15 ± 6	GUTZ	14	DPWA Multichannel	

$\Gamma(N\rho, S=1/2, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17 ± 6	SHRESTHA	12A	DPWA Multichannel	
17 ± 1	VRANA	00	DPWA Multichannel	

*N*(1710) BREIT-WIGNER PHOTON DECAY AMPLITUDES

*N*(1710)  $\rightarrow \rho\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.050 ± 0.010	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.05 ± 0.01	GUTZ	14	DPWA Multichannel
−0.050 ± 0.001	SHKLYAR	13	DPWA Multichannel
0.052 ± 0.015	ANISOVICH	12A	DPWA Multichannel
−0.008 ± 0.003	SHRESTHA	12A	DPWA Multichannel
0.044	PENNER	02D	DPWA Multichannel

*N*(1710)  $\rightarrow n\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.040 ± 0.020	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.017 ± 0.003	SHRESTHA	12A	DPWA Multichannel
−0.024	PENNER	02D	DPWA Multichannel

*N*(1710) FOOTNOTES

- <sup>1</sup> Fit to the amplitudes of HOEHLER 79.  
<sup>2</sup> BATINIC 10 finds evidence for a second  $P_{11}$  state with all parameters except for the phase of the pole residue very similar to the parameters we give here.

*N*(1710) REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
SHKLYAR	05	PR C72 015210	V. Shklyar, H. Lenske, U. Mosel	(GIES)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	90	PR D42 235	R.E. Cutkosky, S. Wang	(CMU)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

*N*(1720)  $3/2^+$

$I(J^P) = \frac{1}{2}(\frac{3}{2}^+)$  Status: \* \* \* \*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

*N*(1720) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1660 to 1690 ( $\approx$ 1675) OUR ESTIMATE			
1670 ± 25	SOKHOYAN	15A	DPWA Multichannel
1677 ± 4 ± 1	SVARC	14	L+P $\pi N \rightarrow \pi N$
1666	ARNDT	06	DPWA $\pi N \rightarrow \pi N$ , $\eta N$
1686	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1680 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1670	SHKLYAR	13	DPWA Multichannel
1660 ± 30	ANISOVICH	12A	DPWA Multichannel
1687	SHRESTHA	12A	DPWA Multichannel
1691 ± 23	BATINIC	10	DPWA $\pi N \rightarrow N\pi$ , $N\eta$
1692	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
150 to 400 ( $\approx$ 250) OUR ESTIMATE			
430 ± 100	SOKHOYAN	15A	DPWA Multichannel
184 ± 8 ± 1	SVARC	14	L+P $\pi N \rightarrow \pi N$
355	ARNDT	06	DPWA $\pi N \rightarrow \pi N$ , $\eta N$
187	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
120 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
118	SHKLYAR	13	DPWA Multichannel
450 ± 100	ANISOVICH	12A	DPWA Multichannel
175	SHRESTHA	12A	DPWA Multichannel
233 ± 23	BATINIC	10	DPWA $\pi N \rightarrow N\pi$ , $N\eta$
94	VRANA	00	DPWA Multichannel





See key on page 601

Baryon Particle Listings

$N(1720)$ ,  $N(1860)$

$\Gamma(N\sigma)/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
8 ± 6	SOKHOYAN 15A	DPWA	Multichannel	
$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<2	SOKHOYAN 15A	DPWA	Multichannel	
$\Gamma(N(1520)\pi, S\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
3 ± 2	SOKHOYAN 15A	DPWA	Multichannel	

$N(1720)$  PHOTON DECAY AMPLITUDES AT THE POLE

$N(1720) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.115 ± 0.045	0 ± 35	SOKHOYAN 15A	DPWA	Multichannel
$N(1720) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.140 ± 0.040	65 ± 35	SOKHOYAN 15A	DPWA	Multichannel

$N(1720)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES

$N(1720) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
0.100 ± 0.020 OUR ESTIMATE				
0.115 ± 0.045	SOKHOYAN 15A	DPWA	Multichannel	
0.095 ± 0.002	WORKMAN 12A	DPWA	$\gamma N \rightarrow N\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.065 ± 0.002	SHKLYAR 13	DPWA	Multichannel	
-0.110 ± 0.045	ANISOVICH 12A	DPWA	Multichannel	
0.057 ± 0.003	SHRESTHA 12A	DPWA	Multichannel	
0.073	DRECHSEL 07	DPWA	$\gamma N \rightarrow \pi N$	
0.097 ± 0.003	DUGGER 07	DPWA	$\gamma N \rightarrow \pi N$	
-0.053	PENNER 02D	DPWA	Multichannel	
$N(1720) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
0.135 ± 0.040	SOKHOYAN 15A	DPWA	Multichannel	
-0.048 ± 0.002	WORKMAN 12A	DPWA	$\gamma N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.035 ± 0.002	SHKLYAR 13	DPWA	Multichannel	
0.150 ± 0.030	ANISOVICH 12A	DPWA	Multichannel	
-0.019 ± 0.002	SHRESTHA 12A	DPWA	Multichannel	
-0.011	DRECHSEL 07	DPWA	$\gamma N \rightarrow \pi N$	
-0.039 ± 0.003	DUGGER 07	DPWA	$\gamma N \rightarrow \pi N$	
0.027	PENNER 02D	DPWA	Multichannel	

$N(1720) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
-0.080 ± 0.050	ANISOVICH 13B	DPWA	Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.002 ± 0.001	SHRESTHA 12A	DPWA	Multichannel	
-0.003	DRECHSEL 07	DPWA	$\gamma N \rightarrow \pi N$	
-0.004	PENNER 02D	DPWA	Multichannel	

$N(1720) \rightarrow n\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
-0.140 ± 0.065	ANISOVICH 13B	DPWA	Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.001 ± 0.002	SHRESTHA 12A	DPWA	Multichannel	
-0.031	DRECHSEL 07	DPWA	$\gamma N \rightarrow \pi N$	
0.003	PENNER 02D	DPWA	Multichannel	

$N(1720)$  FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

$N(1720)$  REFERENCES

For early references, see Physics Letters **111B 1** (1982).

SOKHOYAN 15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG 14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC 14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH 13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR 13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH 12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA 12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN 12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
BATINIC 10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
THOMA 08	PL B659 87	U. Thoma <i>et al.</i>	(CB-ELSA Collab.)

DRECHSEL 07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER 07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT 06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
SHKLYAR 05	PR C72 015210	V. Shklyar, H. Lenske, U. Mosel	(GIES)
PENNER 02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER 02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA 00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER 93	$\pi N$ Newsletter 9 1	G. Hohlner	(KARL)
CUTKOSKY 80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also	PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER 79	PDAT 12-1	G. Hohlner <i>et al.</i>	(KARLT) IJP
Also	Toronto Conf. 3	R. Koch	(KARLT) IJP

$N(1860) 5/2^+$

$I(J^P) = \frac{1}{2}(\frac{5}{2}^+)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

Before the 2012 *Review*, all the evidence for a  $J^P = 5/2^+$  state with a mass above 1800 MeV was filed under a two-star  $N(2000)$ . There is now some evidence from ANISOVICH 12A for two  $5/2^+$  states in this region, so we have split the older data (according to mass) between two two-star  $5/2^+$  states, an  $N(1860)$  and an  $N(2000)$ .

$N(1860)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1834 ± 19 ± 6	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
1830 <sup>+120</sup> - 60	ANISOVICH 12A	DPWA	Multichannel
1807	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1863	SHRESTHA 12A	DPWA	Multichannel

-2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
122 ± 34 ± 7	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
250 <sup>+150</sup> - 50	ANISOVICH 12A	DPWA	Multichannel
109	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
189	SHRESTHA 12A	DPWA	Multichannel

$N(1860)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
4 ± 1 ± 1	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
50 ± 20	ANISOVICH 12A	DPWA	Multichannel
60	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
-39 ± 18 ± 9	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
-80 ± 40	ANISOVICH 12A	DPWA	Multichannel
-67	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$

$N(1860)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1820 to 1960 (≈ 1860) OUR ESTIMATE</b>			
1860 <sup>+120</sup> - 60	ANISOVICH 12A	DPWA	Multichannel
1817.7	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
1882 ± 10	HOEHLER 79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1900 ± 7	SHRESTHA 12A	DPWA	Multichannel

$N(1860)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
270 <sup>+140</sup> - 50	ANISOVICH 12A	DPWA	Multichannel
117.6	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
95 ± 20	HOEHLER 79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
219 ± 23	SHRESTHA 12A	DPWA	Multichannel

# Baryon Particle Listings

## *N*(1860), *N*(1875)

### *N*(1860) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	4–20 %
$\Gamma_2$ $N\eta$	seen
$\Gamma_3$ $N\pi\pi$	
$\Gamma_4$ $N\sigma$	seen
$\Gamma_5$ $p\gamma$	
$\Gamma_6$ $p\gamma$ , helicity=1/2	seen
$\Gamma_7$ $p\gamma$ , helicity=3/2	seen
$\Gamma_8$ $n\gamma$	
$\Gamma_9$ $n\gamma$ , helicity=1/2	
$\Gamma_{10}$ $n\gamma$ , helicity=3/2	

### *N*(1860) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
20 ±6	ANISOVICH	12A	DPWA	Multichannel
12.7	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
4 ±2	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17 ±1	SHRESTHA	12A	DPWA	Multichannel
$\Gamma(N\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4±2	SHRESTHA	12A	DPWA	Multichannel
$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
41±6	SHRESTHA	12A	DPWA	Multichannel

### *N*(1860) BREIT-WIGNER PHOTON DECAY AMPLITUDES

$N(1860) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$	DOCUMENT ID	TECN	COMMENT
VALUE (GeV <sup>-1/2</sup> )			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.017±0.003	SHRESTHA	12A	DPWA Multichannel
$N(1860) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$	DOCUMENT ID	TECN	COMMENT
VALUE			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.029±0.004	SHRESTHA	12A	DPWA Multichannel
$N(1860) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$	DOCUMENT ID	TECN	COMMENT
VALUE (GeV <sup>-1/2</sup> )			
0.021±0.013	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.010±0.005	SHRESTHA	12A	DPWA Multichannel
$N(1860) \rightarrow n\gamma$ , helicity-3/2 amplitude $A_{3/2}$	DOCUMENT ID	TECN	COMMENT
VALUE (GeV <sup>-1/2</sup> )			
0.034±0.017	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.009±0.005	SHRESTHA	12A	DPWA Multichannel

### *N*(1860) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### *N*(1860) REFERENCES

SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT)

## *N*(1875) 3/2<sup>−</sup>

$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$  Status: \*\*\*

Before the 2012 *Review*, all the evidence for a  $J^P = 3/2^-$  state with a mass above 1800 MeV was found under a two-star *N*(2080). There is now evidence from ANISOVICH 12A for two 3/2<sup>−</sup> states in this region, so we have split the older data (according to mass) between a three-star *N*(1875) and a two-star *N*(2120).

### *N*(1875) POLE POSITION

REAL PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
<b>1800 to 1950 OUR ESTIMATE</b>			
1870 ± 20	SOKHOYAN	15A	DPWA Multichannel
2094 ± 7 ± 11	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1880 ± 100	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (lower <i>m</i> )
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1810	SHKLYAR	13	DPWA Multichannel
1860 ± 25	ANISOVICH	12A	DPWA Multichannel
1975	SHRESTHA	12A	DPWA Multichannel
1957 ± 49	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1824	VRANA	00	DPWA Multichannel
−2×IMAGINARY PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
<b>150 to 250 OUR ESTIMATE</b>			
200 ± 15	SOKHOYAN	15A	DPWA Multichannel
296 ± 15 ± 4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
160 ± 80	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (lower <i>m</i> )
• • • We do not use the following data for averages, fits, limits, etc. • • •			
98	SHKLYAR	13	DPWA Multichannel
200 ± 20	ANISOVICH	12A	DPWA Multichannel
495	SHRESTHA	12A	DPWA Multichannel
467 ± 106	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
614	VRANA	00	DPWA Multichannel

### *N*(1875) ELASTIC POLE RESIDUE

MODULUS $ r $	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
<b>2 to 10 OUR ESTIMATE</b>			
3 ± 1.5	SOKHOYAN	15A	DPWA Multichannel
13 ± 1 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
10 ± 5	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (lower <i>m</i> )
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3	SHKLYAR	13	DPWA Multichannel
2.5 ± 1.0	ANISOVICH	12A	DPWA Multichannel
53	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
PHASE $\theta$	DOCUMENT ID	TECN	COMMENT
VALUE (°)			
160 ± 50	SOKHOYAN	15A	DPWA Multichannel
− 2 ± 4 ± 9	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
100 ± 80	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (lower <i>m</i> )
• • • We do not use the following data for averages, fits, limits, etc. • • •			
− 76	SHKLYAR	13	DPWA Multichannel
− 65	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

### *N*(1875) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{\text{pole}}/2$ .

Normalized residue in $N\pi \rightarrow N(1875) \rightarrow \Lambda K$	DOCUMENT ID	TECN	COMMENT
MODULUS			
0.015 ± 0.005	ANISOVICH	12A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(1875) \rightarrow \Sigma K$	DOCUMENT ID	TECN	COMMENT
MODULUS			
0.04 ± 0.02	ANISOVICH	12A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(1875) \rightarrow N\sigma$	DOCUMENT ID	TECN	COMMENT
MODULUS			
0.09 ± 0.03	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.08 ± 0.03	ANISOVICH	12A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(1875) \rightarrow \Delta(1232)\pi$ , S-wave	DOCUMENT ID	TECN	COMMENT
MODULUS			
0.05 ± 0.03	SOKHOYAN	15A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(1875) \rightarrow \Delta(1232)\pi$ , D-wave	DOCUMENT ID	TECN	COMMENT
MODULUS			
0.04 ± 0.02	SOKHOYAN	15A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(1875) \rightarrow N(1440)\pi$	DOCUMENT ID	TECN	COMMENT
MODULUS			
0.03 ± 0.02	SOKHOYAN	15A	DPWA Multichannel

### *N*(1875) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1820 to 1920 (<math>\approx</math> 1875) OUR ESTIMATE</b>			
1875 ± 20	SOKHOYAN	15A	DPWA Multichannel
1934 ± 10	SHKLYAR	13	DPWA Multichannel
1880 ± 100	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

See key on page 601

Baryon Particle Listings  
*N*(1875)

• • • We do not use the following data for averages, fits, limits, etc. • • •

1880 ± 20	ANISOVICH	12A	DPWA	Multichannel
1951 ± 27	SHRESTHA	12A	DPWA	Multichannel
2048 ± 65	BATINIC	10	DPWA	$\pi N \rightarrow N\pi, N\eta$
1946 ± 1	PENNER	02c	DPWA	Multichannel
1895	MART	00	DPWA	$\gamma p \rightarrow \Lambda K^+$
2003 ± 18	VRANA	00	DPWA	Multichannel

*N*(1875) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>250 ± 70 OUR ESTIMATE</b>			
200 ± 25	SOKHOYAN	15A	DPWA Multichannel
857 ± 100	SHKLYAR	13	DPWA Multichannel
180 ± 60	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (lower <i>m</i> )
• • • We do not use the following data for averages, fits, limits, etc. • • •			
200 ± 25	ANISOVICH	12A	DPWA Multichannel
500 ± 45	SHRESTHA	12A	DPWA Multichannel
529 ± 128	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
859 ± 7	PENNER	02c	DPWA Multichannel
372	MART	00	DPWA $\gamma p \rightarrow \Lambda K^+$
1070 ± 858	VRANA	00	DPWA Multichannel

*N*(1875) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ <i>N</i> π	2–14 %
$\Gamma_2$ <i>N</i> η	<1 %
$\Gamma_3$ <i>N</i> ω	15–25 %
$\Gamma_4$ $\Lambda K$	seen
$\Gamma_5$ Σ <i>K</i>	seen
$\Gamma_6$ <i>N</i> ππ	
$\Gamma_7$ Δ(1232)π	10–35 %
$\Gamma_8$ Δ(1232)π, <i>S</i> -wave	7–21 %
$\Gamma_9$ Δ(1232)π, <i>D</i> -wave	2–12 %
$\Gamma_{10}$ <i>N</i> ρ, <i>S</i> =3/2, <i>S</i> -wave	seen
$\Gamma_{11}$ <i>N</i> σ	30–60 %
$\Gamma_{12}$ <i>N</i> (1440)π	2–8 %
$\Gamma_{13}$ <i>N</i> (1520)π	<2 %
$\Gamma_{14}$ <i>p</i> γ	0.001–0.025 %
$\Gamma_{15}$ <i>p</i> γ, helicity=1/2	0.001–0.021 %
$\Gamma_{16}$ <i>p</i> γ, helicity=3/2	<0.003 %
$\Gamma_{17}$ <i>n</i> γ	<0.040 %
$\Gamma_{18}$ <i>n</i> γ, helicity=1/2	<0.007 %
$\Gamma_{19}$ <i>n</i> γ, helicity=3/2	<0.033 %

*N*(1875) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
<b>7 ± 6 OUR ESTIMATE</b>				
4 ± 2	SOKHOYAN	15A	DPWA Multichannel	
11 ± 1	SHKLYAR	13	DPWA Multichannel	
10 ± 4	<sup>2</sup> CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (lower <i>m</i> )	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3 ± 2	ANISOVICH	12A	DPWA Multichannel	
7 ± 2	SHRESTHA	12A	DPWA Multichannel	
17 ± 7	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
12 ± 2	PENNER	02c	DPWA Multichannel	
13 ± 3	VRANA	00	DPWA Multichannel	
$\Gamma(N\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
0 ± 1	SHKLYAR	13	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8 ± 3	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
7 ± 2	PENNER	02c	DPWA Multichannel	
0 ± 2	VRANA	00	DPWA Multichannel	
$\Gamma(N\omega)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
VALUE (%)				
20 ± 5	SHKLYAR	13	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
21 ± 7	PENNER	02c	DPWA Multichannel	
$\Gamma(\Lambda K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.2 ± 0.2	PENNER	02c	DPWA Multichannel	

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE (%)				
0.7 ± 0.4	PENNER	02c	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, S\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE (%)				
14 ± 7	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
87 ± 3	SHRESTHA	12A	DPWA Multichannel	
40 ± 10	VRANA	00	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
VALUE (%)				
7 ± 5	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 6	SHRESTHA	12A	DPWA Multichannel	
17 ± 10	VRANA	00	DPWA Multichannel	

$\Gamma(N\rho, S=3/2, S\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 5	SHRESTHA	12A	DPWA Multichannel	
6 ± 6	VRANA	00	DPWA Multichannel	

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma$
VALUE (%)				
45 ± 15	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 4	SHRESTHA	12A	DPWA Multichannel	
24 ± 24	VRANA	00	DPWA Multichannel	

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12}/\Gamma$
VALUE (%)				
5 ± 3	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1520)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{13}/\Gamma$
VALUE (%)				
< 2	SOKHOYAN	15A	DPWA Multichannel	

*N*(1875) PHOTON DECAY AMPLITUDES AT THE POLE

*N*(1875) → *p*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.017 ± 0.009	−110 ± 40	SOKHOYAN	15A	DPWA Multichannel

*N*(1875) → *p*γ, helicity-3/2 amplitude *A*<sub>3/2</sub>

MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.008 ± 0.004	180 ± 40	SOKHOYAN	15A	DPWA Multichannel

*N*(1875) BREIT-WIGNER PHOTON DECAY AMPLITUDES

*N*(1875) → *p*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.018 ± 0.010	ANISOVICH	12A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.011 ± 0.001	SHKLYAR	13	DPWA Multichannel
0.007 ± 0.008	SHRESTHA	12A	DPWA Multichannel
0.012	PENNER	02D	DPWA Multichannel

*N*(1875) → *p*γ, helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.007 ± 0.004	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.026 ± 0.001	SHKLYAR	13	DPWA Multichannel
−0.009 ± 0.005	ANISOVICH	12A	DPWA Multichannel
0.043 ± 0.022	SHRESTHA	12A	DPWA Multichannel
−0.010	PENNER	02D	DPWA Multichannel

*N*(1875) → *n*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.010 ± 0.006	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.055 ± 0.021	SHRESTHA	12A	DPWA Multichannel
0.023	PENNER	02D	DPWA Multichannel

# Baryon Particle Listings

## $N(1875)$ , $N(1880)$

### $N(1875) \rightarrow n\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.020±0.015	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
−0.085±0.031	SHRESTHA	12A	DPWA Multichannel
−0.009	PENNER	02D	DPWA Multichannel

### N(1875) FOOTNOTES

- <sup>1</sup> Fit to the amplitudes of HOEHLER 79.  
<sup>2</sup> CUTKOSKY 80 finds a lower mass  $D_{13}$  resonance, as well as one in this region. Both are listed here.

### N(1875) REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
MART	00	PR C61 012201	T. Mart, C. Bennhold	
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. H�hler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$N(1880) \ 1/2^{+}$

$I(J^P) = \frac{1}{2}(\frac{1}{2}^{+})$  Status: \*\*

OMITTED FROM SUMMARY TABLE

### N(1880) POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1870±40	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
1870±40	GUTZ	14	DPWA Multichannel
1860±35	ANISOVICH	12A	DPWA Multichannel
1801	SHRESTHA	12A	DPWA Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
220±50	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
220±50	GUTZ	14	DPWA Multichannel
250±70	ANISOVICH	12A	DPWA Multichannel
383	SHRESTHA	12A	DPWA Multichannel

### N(1880) ELASTIC POLE RESIDUE

#### MODULUS $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
6±4	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
6±4	GUTZ	14	DPWA Multichannel
6±4	ANISOVICH	12A	DPWA Multichannel

#### PHASE $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
70±60	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
70±60	GUTZ	14	DPWA Multichannel
80±65	ANISOVICH	12A	DPWA Multichannel

### N(1880) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow N\eta$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.11±0.07	−75 ± 55	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow \Lambda K$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.03±0.02	40 ± 40	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow \Sigma K$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.11±0.06	95 ± 40	ANISOVICH	12A	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow \Delta\pi$ , $P$ -wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.14±0.08	−150 ± 55	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
0.20±0.08	−150 ± 50	ANISOVICH	12A	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow N(1535)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.09±0.05	130 ± 60	GUTZ	14	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow N a_0(980)$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.04±0.03	40 ± 65	GUTZ	14	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow N(1880) \rightarrow N\sigma$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.10±0.05	−140 ± 55	SOKHOYAN	15A	DPWA Multichannel

### N(1880) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1875±40	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
1875±40	GUTZ	14	DPWA Multichannel
1870±35	ANISOVICH	12A	DPWA Multichannel
1900±36	SHRESTHA	12A	DPWA Multichannel

### N(1880) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
230± 50	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
230± 50	GUTZ	14	DPWA Multichannel
235± 65	ANISOVICH	12A	DPWA Multichannel
485±142	SHRESTHA	12A	DPWA Multichannel

### N(1880) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \ N\pi$	3–9 %
$\Gamma_2 \ N\eta$	5–55 %
$\Gamma_3 \ \Lambda K$	1–3 %
$\Gamma_4 \ \Sigma K$	10–24 %
$\Gamma_5 \ N\pi\pi$	30–80 %
$\Gamma_6 \ \Delta(1232)\pi$	18–42 %
$\Gamma_7 \ N\sigma$	10–40 %
$\Gamma_8 \ N(1535)\pi$	4–12 %
$\Gamma_9 \ N a_0(980)$	1–5 %
$\Gamma_{10} \ p\gamma$ , helicity=1/2	seen
$\Gamma_{11} \ n\gamma$ , helicity=1/2	0.002–0.63 %

### N(1880) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
6 ± 3	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6 ± 3	GUTZ	14	DPWA	Multichannel
5 ± 3	ANISOVICH	12A	DPWA	Multichannel
15 ± 5	SHRESTHA	12A	DPWA	Multichannel

$\Gamma(N\eta)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$25^{+30}_{-20}$	ANISOVICH	12A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$16 \pm 7$	SHRESTHA	12A	DPWA Multichannel

$\Gamma(\Lambda K)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$2 \pm 1$	ANISOVICH	12A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$32 \pm 10$	SHRESTHA	12A	DPWA Multichannel

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
17±7	ANISOVICH	12A	DPWA Multichannel

See key on page 601

Baryon Particle Listings  
*N*(1880), *N*(1895)

$\Gamma(\Delta(1232)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)			
30 ± 12	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
29 ± 12	ANISOVICH	12A	DPWA Multichannel
< 2	SHRESTHA	12A	DPWA Multichannel

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)			
25 ± 15	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
8 ± 5	SHRESTHA	12A	DPWA Multichannel

$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)			
8 ± 4	GUTZ	14	DPWA Multichannel

$\Gamma(Na_0(980))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)			
3 ± 2	GUTZ	14	DPWA Multichannel

*N*(1880) BREIT-WIGNER PHOTON DECAY AMPLITUDES

<i>N</i> (1880) → <i>p</i> γ, helicity-1/2 amplitude <i>A</i> <sub>1/2</sub>	DOCUMENT ID	TECN	COMMENT
VALUE (GeV <sup>-1/2</sup> )			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.021 ± 0.006	SHRESTHA	12A	DPWA Multichannel

<i>N</i> (1880) → <i>n</i> γ, helicity-1/2 amplitude <i>A</i> <sub>1/2</sub>	DOCUMENT ID	TECN	COMMENT
VALUE (GeV <sup>-1/2</sup> )			
− 0.060 ± 0.050	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.014 ± 0.007	SHRESTHA	12A	DPWA Multichannel

*N*(1880) REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)

*N*(1895) 1/2<sup>−</sup>

$I(J^P) = \frac{1}{2}(\frac{1}{2}^-)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

Before our 2012 *Review*, this state appeared in our Listings as the *N*(2090). Any structure in the *S*<sub>11</sub> wave above 1800 MeV is listed here. A few early results that are now obsolete have been omitted.

*N*(1895) POLE POSITION

REAL PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
1907 ± 10	SOKHOYAN	15A	DPWA Multichannel
1917 ± 19 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2150 ± 70	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1900 ± 15	ANISOVICH	12A	DPWA Multichannel
1858	SHRESTHA	12A	DPWA Multichannel
1797 ± 26	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1795	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
100 <sup>+</sup> <sub>−15</sub>	SOKHOYAN	15A	DPWA Multichannel
101 ± 36 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
350 ± 100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
90 <sup>+</sup> <sub>−15</sub>	ANISOVICH	12A	DPWA Multichannel
479	SHRESTHA	12A	DPWA Multichannel
420 ± 45	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
220	VRANA	00	DPWA Multichannel

*N*(1895) ELASTIC POLE RESIDUE

MODULUS   <i>r</i>	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
3 ± 2	SOKHOYAN	15A	DPWA Multichannel
3.1 ± 1.4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
40 ± 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1 ± 1	ANISOVICH	12A	DPWA Multichannel
60	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

PHASE θ	DOCUMENT ID	TECN	COMMENT
VALUE (°)			
125 ± 45	SOKHOYAN	15A	DPWA Multichannel
−107 ± 23 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
0 ± 90	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−164	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

*N*(1895) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{\text{pole}}/2$ .

Normalized residue in <i>N</i> π → <i>N</i> (1895) → <i>N</i> η	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE (°)		
0.06 ± 0.02	40 ± 20	ANISOVICH	12A DPWA Multichannel

Normalized residue in <i>N</i> π → <i>N</i> (1895) → Λ <i>K</i>	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE (°)		
0.05 ± 0.02	− 90 ± 30	ANISOVICH	12A DPWA Multichannel

Normalized residue in <i>N</i> π → <i>N</i> (1895) → Σ <i>K</i>	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE (°)		
0.06 ± 0.02	40 ± 30	ANISOVICH	12A DPWA Multichannel

Normalized residue in <i>N</i> π → <i>N</i> (1895) → Δ(1232)π	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE (°)		
0.05 ± 0.025	− 100 ± 45	SOKHOYAN	15A DPWA Multichannel

Normalized residue in <i>N</i> π → <i>N</i> (1895) → <i>N</i> (1440)π	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE (°)		
0.05 ± 0.025	− 100 ± 45	SOKHOYAN	15A DPWA Multichannel

*N*(1895) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1905 ± 12	SOKHOYAN	15A	DPWA Multichannel
2180 ± 80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1880 ± 20	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1895 ± 15	ANISOVICH	12A	DPWA Multichannel
1910 ± 15	SHRESTHA	12A	DPWA Multichannel
1812 ± 25	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1822 ± 43	VRANA	00	DPWA Multichannel

*N*(1895) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
100 <sup>+</sup> <sub>−30</sub>	SOKHOYAN	15A	DPWA Multichannel
350 ± 100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
95 ± 30	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
90 <sup>+</sup> <sub>−15</sub>	ANISOVICH	12A	DPWA Multichannel
502 ± 47	SHRESTHA	12A	DPWA Multichannel
405 ± 40	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
248 ± 185	VRANA	00	DPWA Multichannel

*N*(1895) DECAY MODES

Mode	Fraction (Γ <sub><i>i</i></sub> /Γ)
Γ <sub>1</sub> <i>N</i> π	1–4 %
Γ <sub>2</sub> <i>N</i> η	15–27 %
Γ <sub>3</sub> Λ <i>K</i>	13–23 %
Γ <sub>4</sub> Σ <i>K</i>	6–20 %
Γ <sub>5</sub> <i>N</i> ππ	
Γ <sub>6</sub> Δ(1232)π	
Γ <sub>7</sub> Δ(1232)π, <i>D</i> -wave	3–11 %
Γ <sub>8</sub> <i>N</i> ρ	
Γ <sub>9</sub> <i>N</i> ρ, <i>S</i> =1/2, <i>S</i> -wave	seen
Γ <sub>10</sub> <i>N</i> ρ, <i>S</i> =3/2, <i>D</i> -wave	seen
Γ <sub>11</sub> <i>N</i> σ	seen
Γ <sub>12</sub> <i>N</i> (1440)π	1–4 %
Γ <sub>13</sub> <i>p</i> γ, helicity=1/2	0.01–0.06 %
Γ <sub>14</sub> <i>n</i> γ, helicity=1/2	0.003–0.05 %

# Baryon Particle Listings

## N(1895), N(1900)

N(1895) BRANCHING RATIOS				
$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
2.5±1.5	SOKHOYAN	15A	DPWA Multichannel	
18 ±8	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
9 ±5	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2 ±1	ANISOVICH	12A	DPWA Multichannel	
17 ±2	SHRESTHA	12A	DPWA Multichannel	
32 ±6	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
17 ±3	VRANA	00	DPWA Multichannel	
$\Gamma(N\eta)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
21± 6	ANISOVICH	12A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
40± 4	SHRESTHA	12A	DPWA Multichannel	
22±10	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$	
41± 4	VRANA	00	DPWA Multichannel	
$\Gamma(\Lambda K)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
18 ±5	ANISOVICH	12A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.8±0.8	SHRESTHA	12A	DPWA Multichannel	
$\Gamma(\Sigma K)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
13±7	ANISOVICH	12A	DPWA Multichannel	
$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
7±4	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7±3	SHRESTHA	12A	DPWA Multichannel	
1±1	VRANA	00	DPWA Multichannel	
$\Gamma(N\rho, S=1/2, S\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_9/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 2	SHRESTHA	12A	DPWA Multichannel	
36±1	VRANA	00	DPWA Multichannel	
$\Gamma(N\rho, S=3/2, D\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_{10}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
9±3	SHRESTHA	12A	DPWA Multichannel	
1±1	VRANA	00	DPWA Multichannel	
$\Gamma(N\sigma)/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<2	SHRESTHA	12A	DPWA Multichannel	
2±1	VRANA	00	DPWA Multichannel	
$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
2.5±1.5	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
24 ±4	SHRESTHA	12A	DPWA Multichannel	
2 ±1	VRANA	00	DPWA Multichannel	

### N(1895) PHOTON DECAY AMPLITUDES AT THE POLE

#### N(1895) → pγ, helicity-1/2 amplitude A<sub>1/2</sub>

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.015±0.006	145 ± 35	SOKHOYAN	15A	DPWA Multichannel

### N(1895) BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### N(1895) → pγ, helicity-1/2 amplitude A<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.016±0.006	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.012±0.006	SHRESTHA	12A	DPWA Multichannel

#### N(1895) → nγ, helicity-1/2 amplitude A<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.013±0.006	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.003±0.007	SHRESTHA	12A	DPWA Multichannel

### N(1895) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### N(1895) REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)JIP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT)JIP
Also		Toronto Conf. 3	R. Koch	(KARLT)JIP

N(1900) 3/2<sup>+</sup>

I(J<sup>P</sup>) =  $\frac{1}{2}(3/2^+)$  Status: \*\*\*

### N(1900) POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1900 to 1940 (≈ 1920) OUR ESTIMATE			
1910±30	SOKHOYAN	15A	DPWA Multichannel
1928±18±2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1910±30	GUTZ	14	DPWA Multichannel
1910	SHKLYAR	13	DPWA Multichannel
1900±30	ANISOVICH	12A	DPWA Multichannel
1895	SHRESTHA	12A	DPWA Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
130 to 300 OUR ESTIMATE			
280± 50	SOKHOYAN	15A	DPWA Multichannel
152± 40±9	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
280± 50	GUTZ	14	DPWA Multichannel
173	SHKLYAR	13	DPWA Multichannel
200+100 − 60	ANISOVICH	12A	DPWA Multichannel
100	SHRESTHA	12A	DPWA Multichannel

### N(1900) ELASTIC POLE RESIDUE

#### MODULUS |r|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
4±2	SOKHOYAN	15A	DPWA Multichannel
4±1±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4±2	GUTZ	14	DPWA Multichannel
10	SHKLYAR	13	DPWA Multichannel
3±2	ANISOVICH	12A	DPWA Multichannel

#### PHASE θ

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−10±40	SOKHOYAN	15A	DPWA Multichannel
−29±15±2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−10±40	GUTZ	14	DPWA Multichannel
−64	SHKLYAR	13	DPWA Multichannel
10±35	ANISOVICH	12A	DPWA Multichannel

### N(1900) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{\text{pole}}/2$ .

#### Normalized residue in Nπ → N(1900) → Nη

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.05±0.02	70 ± 60	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in Nπ → N(1900) → ΛK

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.07±0.03	135 ± 25	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in Nπ → N(1900) → ΣK

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.04±0.02	110 ± 30	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in Nπ → N(1900) → N(1535)π

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.04±0.01	170 ± 30	GUTZ	14	DPWA Multichannel

See key on page 601

# Baryon Particle Listings

## *N*(1900)

Normalized residue in $N\pi \rightarrow N(1900) \rightarrow \Delta(1232)\pi$ , <i>P</i> -wave				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.07±0.04	−65 ± 30	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in $N\pi \rightarrow N(1900) \rightarrow \Delta(1232)\pi$ , <i>F</i> -wave				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.10±0.05	80 ± 30	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in $N\pi \rightarrow N(1900) \rightarrow N(1520)\pi$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.07±0.04	−105 ± 35	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in $N\pi \rightarrow N(1900) \rightarrow N\sigma$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.03±0.02	−110 ± 35	SOKHOYAN	15A	DPWA Multichannel

### *N*(1900) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1900±30 OUR ESTIMATE</b>			
1910±30	SOKHOYAN	15A	DPWA Multichannel
1998 ± 3	SHKLYAR	13	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1910±30	GUTZ	14	DPWA Multichannel
1905 ± 30	ANISOVICH	12A	DPWA Multichannel
1900 ± 8	SHRESTHA	12A	DPWA Multichannel
1951±53	PENNER	02C	DPWA Multichannel

### *N*(1900) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>200± 50 OUR ESTIMATE</b>			
270± 50	SOKHOYAN	15A	DPWA Multichannel
359± 10	SHKLYAR	13	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
270± 50	GUTZ	14	DPWA Multichannel
250+120 − 50	ANISOVICH	12A	DPWA Multichannel
101± 15	SHRESTHA	12A	DPWA Multichannel
622± 42	PENNER	02C	DPWA Multichannel

### *N*(1900) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	<10 %
$\Gamma_2$ $N\eta$	2–14 %
$\Gamma_3$ $N\omega$	7–13 %
$\Gamma_4$ $\Lambda K$	2–20 %
$\Gamma_5$ $\Sigma K$	3–7 %
$\Gamma_6$ $N\pi\pi$	40–80 %
$\Gamma_7$ $\Delta(1232)\pi$	30–70 %
$\Gamma_8$ $\Delta(1232)\pi$ , <i>P</i> -wave	9–25 %
$\Gamma_9$ $\Delta(1232)\pi$ , <i>F</i> -wave	21–45 %
$\Gamma_{10}$ $N\sigma$	1–7 %
$\Gamma_{11}$ $N(1520)\pi$	7–23 %
$\Gamma_{12}$ $N(1535)\pi$	4–10 %
$\Gamma_{13}$ $p\gamma$	0.001–0.025 %
$\Gamma_{14}$ $p\gamma$ , helicity=1/2	0.001–0.021 %
$\Gamma_{15}$ $p\gamma$ , helicity=3/2	<0.003 %
$\Gamma_{16}$ $n\gamma$	<0.040 %
$\Gamma_{17}$ $n\gamma$ , helicity=1/2	<0.007 %
$\Gamma_{18}$ $n\gamma$ , helicity=3/2	<0.033 %

### *N*(1900) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
3±2	SOKHOYAN	15A	DPWA Multichannel	
25±1	SHKLYAR	13	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3±2	GUTZ	14	DPWA Multichannel	
3±2	ANISOVICH	12A	DPWA Multichannel	
7±4	SHRESTHA	12A	DPWA Multichannel	
16±2	PENNER	02C	DPWA Multichannel	

$\Gamma(N\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
2±2	SHKLYAR	13	DPWA Multichannel	
10±4	ANISOVICH	12A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 1	SHRESTHA	12A	DPWA Multichannel	
14±5	PENNER	02C	DPWA Multichannel	

$\Gamma(N\omega)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
VALUE (%)				
10±3	SHKLYAR	13	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
39±9	PENNER	02C	DPWA Multichannel	

$\Gamma(\Lambda K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE (%)				
16 ± 5	ANISOVICH	12A	DPWA Multichannel	
2.4±0.3	SHKLYAR	05	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
14 ± 5	SHRESTHA	12A	DPWA Multichannel	
5 to 15	NIKONOV	08	DPWA Multichannel	
0.1±0.1	PENNER	02C	DPWA Multichannel	

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE (%)				
5±2	ANISOVICH	12A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1±1	PENNER	02C	DPWA Multichannel	

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
VALUE (%)				
4±3	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1520)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma$
VALUE (%)				
15±8	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12}/\Gamma$
VALUE (%)				
7±3	GUTZ	14	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE (%)				
17±8	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
VALUE (%)				
33±12	SOKHOYAN	15A	DPWA Multichannel	

### *N*(1900) PHOTON DECAY AMPLITUDES AT THE POLE

$N(1900) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.026±0.014	60 ± 35	SOKHOYAN	15A	DPWA Multichannel

$N(1900) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
−0.070±0.030	70 ± 50	SOKHOYAN	15A	DPWA Multichannel

### *N*(1900) BREIT-WIGNER PHOTON DECAY AMPLITUDES

$N(1900) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
0.024±0.014	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.024±0.014	GUTZ	14	DPWA Multichannel	
−0.008±0.001	SHKLYAR	13	DPWA Multichannel	
0.026±0.015	ANISOVICH	12A	DPWA Multichannel	
0.041±0.008	SHRESTHA	12A	DPWA Multichannel	
−0.017	PENNER	02D	DPWA Multichannel	

$N(1900) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
−0.067±0.030	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.067±0.030	GUTZ	14	DPWA Multichannel	
0. ±0.001	SHKLYAR	13	DPWA Multichannel	
−0.065±0.030	ANISOVICH	12A	DPWA Multichannel	
−0.004±0.006	SHRESTHA	12A	DPWA Multichannel	
0.031	PENNER	02D	DPWA Multichannel	

Baryon Particle Listings

*N*(1900), *N*(1990), *N*(2000)

*N*(1900) → *n*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.000 ± 0.030	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.010 ± 0.004	SHRESTHA	12A	DPWA Multichannel
−0.016	PENNER	02D	DPWA Multichannel

*N*(1900) → *n*γ, helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.060 ± 0.045	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.011 ± 0.007	SHRESTHA	12A	DPWA Multichannel
−0.002	PENNER	02D	DPWA Multichannel

*N*(1900) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

*N*(1900) REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosel	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
NIKONOV	08	PL B662 245	V.A. Nikonov <i>et al.</i>	(Bonn, Gatchina)
SHKLYAR	05	PR C72 015210	V. Shklyar, H. Lenske, U. Mosel	(GIES)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT)

*N*(1990) 7/2<sup>+</sup>

*I*(*J*<sup>*P*</sup>) =  $\frac{1}{2}(\frac{7}{2}^+)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

*N*(1990) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2030 ± 65	ANISOVICH	12A	DPWA Multichannel
1900 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1941	SHRESTHA	12A	DPWA Multichannel
2301	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
240 ± 60	ANISOVICH	12A	DPWA Multichannel
260 ± 60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
130	SHRESTHA	12A	DPWA Multichannel
202	VRANA	00	DPWA Multichannel

*N*(1990) ELASTIC POLE RESIDUE

MODULUS |*r*|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2 ± 1	ANISOVICH	12A	DPWA Multichannel
9 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

PHASE θ

VALUE (°)	DOCUMENT ID	TECN	COMMENT
125 ± 65	ANISOVICH	12A	DPWA Multichannel
− 60 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

*N*(1990) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2060 ± 65	ANISOVICH	12A	DPWA Multichannel
1970 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2005 ± 150	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1990 ± 45	SHRESTHA	12A	DPWA Multichannel
2311 ± 16	VRANA	00	DPWA Multichannel

*N*(1990) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>200 to 400 (≈ 300) OUR ESTIMATE</b>			
240 ± 50	ANISOVICH	12A	DPWA Multichannel
350 ± 120	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
350 ± 100	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			

203 ± 161	SHRESTHA	12A	DPWA Multichannel
205 ± 72	VRANA	00	DPWA Multichannel

*N*(1990) DECAY MODES

Mode	Fraction (Γ <sub><i>i</i></sub> /Γ)
Γ <sub>1</sub> <i>N</i> π	2–6 %
Γ <sub>2</sub> <i>p</i> γ	0.01–0.12 %
Γ <sub>3</sub> <i>p</i> γ, helicity=1/2	0.003–0.042 %
Γ <sub>4</sub> <i>p</i> γ, helicity=3/2	0.009–0.075 %
Γ <sub>5</sub> <i>n</i> γ	0.01–0.16 %
Γ <sub>6</sub> <i>n</i> γ, helicity=1/2	0.003–0.066 %
Γ <sub>7</sub> <i>n</i> γ, helicity=3/2	0.003–0.098 %

*N*(1990) BRANCHING RATIOS

Γ( <i>N</i> π)/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT	Γ <sub>1</sub> /Γ
<b>2 to 6 (≈ 4) OUR ESTIMATE</b>				
2 ± 1	ANISOVICH	12A	DPWA Multichannel	
6 ± 2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
4 ± 2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2 ± 1	SHRESTHA	12A	DPWA Multichannel	
22 ± 11	VRANA	00	DPWA Multichannel	

*N*(1990) BREIT-WIGNER PHOTON DECAY AMPLITUDES

*N*(1990) → *p*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.040 ± 0.012	ANISOVICH	12A	DPWA Multichannel

*N*(1990) → *p*γ, helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.057 ± 0.012	ANISOVICH	12A	DPWA Multichannel

*N*(1990) → *n*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.045 ± 0.020	ANISOVICH	13B	DPWA Multichannel

*N*(1990) → *n*γ, helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
−0.052 ± 0.027	ANISOVICH	13B	DPWA Multichannel

*N*(1990) REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

*N*(2000) 5/2<sup>+</sup>

*I*(*J*<sup>*P*</sup>) =  $\frac{1}{2}(\frac{5}{2}^+)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

Before the 2012 *Review*, all the evidence for a *J*<sup>*P*</sup> = 5/2<sup>+</sup> state with a mass above 1800 MeV was filed under a two-star *N*(2000). There is now some evidence from ANISOVICH 12A for two 5/2<sup>+</sup> states in this region, so we have split the older data (according to mass) between two two-star 5/2<sup>+</sup> states, an *N*(1860) and an *N*(2000).

*N*(2000) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2030 ± 40	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1900	SHKLYAR	13	DPWA Multichannel
2030 ± 110	ANISOVICH	12A	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
380 ± 60	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
123	SHKLYAR	13	DPWA Multichannel
480 ± 100	ANISOVICH	12A	DPWA Multichannel



See key on page 601

Baryon Particle Listings

$N(2000)$ ,  $N(2040)$

$N(2000)$ ELASTIC POLE RESIDUE				
MODULUS $ r $				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
18± 8	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
11	SHKLYAR	13	DPWA	Multichannel
35 <sup>+80</sup> <sub>-15</sub>	ANISOVICH	12A	DPWA	Multichannel
PHASE $\theta$				
VALUE (°)	DOCUMENT ID	TECN	COMMENT	
-150±40	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
- 6	SHKLYAR	13	DPWA	Multichannel
-100±40	ANISOVICH	12A	DPWA	Multichannel

$N(2000)$ INELASTIC POLE RESIDUE				
The "normalized residue" is the residue divided by $\Gamma_{pole}/2$ .				
Normalized residue in $N\pi \rightarrow N(2000) \rightarrow \Delta(1232)\pi$ , $P$ -wave				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.16±0.06	100 ± 50	SOKHOYAN	15A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(2000) \rightarrow \Delta(1232)\pi$ , $F$ -wave				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.20±0.10	- 20 ± 45	SOKHOYAN	15A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(2000) \rightarrow N\sigma$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.12±0.06	80 ± 40	SOKHOYAN	15A	DPWA Multichannel
Normalized residue in $N\pi \rightarrow N(2000) \rightarrow N(1520)\pi$ , $D$ -wave				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.17±0.09	- 60 ± 35	SOKHOYAN	15A	DPWA Multichannel

$N(2000)$ BREIT-WIGNER MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
2060± 30	SOKHOYAN	15A	DPWA	Multichannel
1946± 4	SHKLYAR	13	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2090±120	ANISOVICH	12A	DPWA	Multichannel

$N(2000)$ BREIT-WIGNER WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
390± 55	SOKHOYAN	15A	DPWA	Multichannel
198± 2	SHKLYAR	13	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
460±100	ANISOVICH	12A	DPWA	Multichannel

$N(2000)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\pi$	6-10 %	
$\Gamma_2$ $N\eta$	<4 %	
$\Gamma_3$ $N\omega$	<2 %	
$\Gamma_4$ $N\pi\pi$	35-90 %	
$\Gamma_5$ $\Delta(1232)\pi$	30-80 %	
$\Gamma_6$ $\Delta(1232)\pi$ , $P$ -wave	12-32 %	
$\Gamma_7$ $\Delta(1232)\pi$ , $F$ -wave	19-49 %	
$\Gamma_8$ $N\sigma$	5-15 %	
$\Gamma_9$ $N(1520)\pi$ , $D$ -wave	11-31 %	
$\Gamma_{10}$ $N(1680)\pi$ , $P$ -wave	17-25 %	
$\Gamma_{11}$ $p\gamma$	0.01-0.08 %	
$\Gamma_{12}$ $p\gamma$ , helicity=1/2	0.003-0.031 %	
$\Gamma_{13}$ $p\gamma$ , helicity=3/2	0.008-0.048 %	
$\Gamma_{14}$ $n\gamma$	0.002-0.07 %	
$\Gamma_{15}$ $n\gamma$ , helicity=1/2	<0.017 %	
$\Gamma_{16}$ $n\gamma$ , helicity=3/2	0.001-0.056 %	

$N(2000)$ BRANCHING RATIOS				
$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma_2$
<u>VALUE (%)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>6 to 10 (<math>\approx</math> 8) OUR ESTIMATE</b>				
8 $\pm$ 4	SOKHOYAN	15A	DPWA	Multichannel
10 $\pm$ 1	SHKLYAR	13	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
9 $\pm$ 4	ANISOVICH	12A	DPWA	Multichannel

$\Gamma(N\eta)/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
2±2	SHKLYAR	13	DPWA	Multichannel
$\Gamma(N\omega)/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_3/\Gamma$
VALUE (%)				
1±1	SHKLYAR	13	DPWA	Multichannel
$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
VALUE (%)				
22±10	SOKHOYAN	15A	DPWA	Multichannel
$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE (%)				
34±15	SOKHOYAN	15A	DPWA	Multichannel
$\Gamma(N\sigma)/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE (%)				
10±5	SOKHOYAN	15A	DPWA	Multichannel
$\Gamma(N(1520)\pi, D\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
VALUE (%)				
21±10	SOKHOYAN	15A	DPWA	Multichannel
$\Gamma(N(1680)\pi, P\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
VALUE (%)				
16±9	SOKHOYAN	15A	DPWA	Multichannel

$N(2000)$ PHOTON DECAY AMPLITUDES AT THE POLE				
$N(2000) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.033±0.010	15 ± 25	SOKHOYAN	15A	DPWA Multichannel
$N(2000) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.045±0.008	-140 ± 25	SOKHOYAN	15A	DPWA Multichannel

$N(2000)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES				
$N(2000) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT	
0.031±0.010	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.011±0.001	SHKLYAR	13	DPWA	Multichannel
$N(2000) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT	
- 0.043±0.008	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.025±0.001	SHKLYAR	13	DPWA	Multichannel
$N(2000) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT	
- 0.018±0.012	ANISOVICH	13B	DPWA	Multichannel
$N(2000) \rightarrow n\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT	
- 0.035±0.020	ANISOVICH	13B	DPWA	Multichannel

$N(2000)$ REFERENCES				
SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan et al.	(CBELSA/TAPS Collab.)
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich et al.	
SHKLYAR	13	PR C87 015201	V. Shklyar, H. Lenske, U. Mosele	(GIES)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich et al.	(BONN, PNPI)

<div><math>N(2040)</math> 3/2<sup>+</sup></div>	$J^P = \frac{3}{2}^+$	Status: *
OMITTED FROM SUMMARY TABLE		

$N(2040)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
2040 <sup>+3</sup> <sub>-4</sub> ± 25	ABLIKIM	09B	BES2	$J/\psi \rightarrow p\bar{p}\pi^0$
2068±3 <sup>+15</sup> <sub>-40</sub>	ABLIKIM	06K	BES2	$J/\psi \rightarrow p\bar{n}\pi^-, n\bar{p}\pi^+$

Baryon Particle Listings

*N*(2040), *N*(2060)

*N*(2040) WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
230 ± 8 ± 52	ABLIKIM	09B BES2	$J/\psi \rightarrow p\overline{p}\pi^0$
165 ± 14 ± 40	ABLIKIM	06K BES2	$J/\psi \rightarrow p\overline{p}\pi^-, n\overline{p}\pi^+$

*N*(2040) REFERENCES

ABLIKIM	09B	PR D80 052004	M. Ablikim <i>et al.</i>	(BES II Collab.)
ABLIKIM	06K	PRL 97 062001	M. Ablikim <i>et al.</i>	(BES II Collab.)

*N*(2060) 5/2<sup>−</sup>

$I(J^P) = \frac{1}{2}(\frac{5}{2}^-)$  Status: \* \*

OMITTED FROM SUMMARY TABLE

Before our 2012 *Review*, this state appeared in our Listings as the *N*(2200).

*N*(2060) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2030 ± 15	SOKHOYAN	15A DPWA	Multichannel
2119 ± 11 ± 1	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
2100 ± 60	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2040 ± 15	ANISOVICH	12A DPWA	Multichannel
2064	SHRESTHA	12A DPWA	Multichannel
2144 ± 31	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
400 ± 35	SOKHOYAN	15A DPWA	Multichannel
370 ± 20 ± 5	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
360 ± 80	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
390 ± 25	ANISOVICH	12A DPWA	Multichannel
267	SHRESTHA	12A DPWA	Multichannel
438 ± 13	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$

*N*(2060) ELASTIC POLE RESIDUE

MODULUS |*r*|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
25 ± 8	SOKHOYAN	15A DPWA	Multichannel
19 ± 1 ± 1	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
20 ± 10	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
19 ± 5	ANISOVICH	12A DPWA	Multichannel
26	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−130 ± 20	SOKHOYAN	15A DPWA	Multichannel
− 94 ± 5 ± 1	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
− 90 ± 50	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−125 ± 20	ANISOVICH	12A DPWA	Multichannel
− 71	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$

*N*(2060) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

Normalized residue in *N*π → *N*(2060) → *N*η

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.05 ± 0.03	40 ± 25	ANISOVICH	12A DPWA	Multichannel

Normalized residue in *N*π → *N*(2060) → Λ*K*

MODULUS	DOCUMENT ID	TECN	COMMENT
0.01 ± 0.005	ANISOVICH	12A DPWA	Multichannel

Normalized residue in *N*π → *N*(2060) → Σ*K*

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.04 ± 0.02	−70 ± 30	ANISOVICH	12A DPWA	Multichannel

Normalized residue in *N*π → *N*(2060) → Δ(1232)π, *D*-wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.06 ± 0.03	−90 ± 40	SOKHOYAN	15A DPWA	Multichannel

Normalized residue in *N*π → *N*(2060) → *N*σ

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.12 ± 0.06	80 ± 40	SOKHOYAN	15A DPWA	Multichannel

Normalized residue in *N*π → *N*(2060) → *N*(1440)π

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.17 ± 0.09	−60 ± 35	SOKHOYAN	15A DPWA	Multichannel

Normalized residue in *N*π → *N*(2060) → *N*(1520)π, *P*-wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.14 ± 0.06	−45 ± 15	SOKHOYAN	15A DPWA	Multichannel

*N*(2060) BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2045 ± 15	SOKHOYAN	15A DPWA	Multichannel
2180 ± 80	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$
2228 ± 30	HOEHLER	79 IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2060 ± 15	ANISOVICH	12A DPWA	Multichannel
2116 ± 21	SHRESTHA	12A DPWA	Multichannel
2217 ± 27	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$

*N*(2060) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
420 ± 30	SOKHOYAN	15A DPWA	Multichannel
400 ± 100	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$
310 ± 50	HOEHLER	79 IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
375 ± 25	ANISOVICH	12A DPWA	Multichannel
307 ± 112	SHRESTHA	12A DPWA	Multichannel
481 ± 17	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$

*N*(2060) DECAY MODES

Mode	Fraction (Γ <sub><i>i</i></sub> /Γ)
Γ <sub>1</sub> <i>N</i> π	7–12 %
Γ <sub>2</sub> <i>N</i> η	2–6 %
Γ <sub>3</sub> Λ <i>K</i>	seen
Γ <sub>4</sub> Σ <i>K</i>	1–5 %
Γ <sub>5</sub> <i>N</i> ππ	
Γ <sub>6</sub> Δ(1232)π	
Γ <sub>7</sub> Δ(1232)π, <i>D</i> -wave	4–10 %
Γ <sub>8</sub> <i>N</i> ρ	
Γ <sub>9</sub> <i>N</i> ρ, <i>S</i> =1/2, <i>P</i> -wave	seen
Γ <sub>10</sub> <i>N</i> σ	3–9 %
Γ <sub>11</sub> <i>N</i> (1440)π	4–14 %
Γ <sub>12</sub> <i>N</i> (1520)π, <i>P</i> -wave	9–21 %
Γ <sub>13</sub> <i>N</i> (1680)π, <i>S</i> -wave	8–22 %
Γ <sub>14</sub> <i>p</i> γ	0.03–0.19 %
Γ <sub>15</sub> <i>p</i> γ, helicity=1/2	0.02–0.08 %
Γ <sub>16</sub> <i>p</i> γ, helicity=3/2	0.01–0.10 %
Γ <sub>17</sub> <i>n</i> γ	0.003–0.07 %
Γ <sub>18</sub> <i>n</i> γ, helicity=1/2	0.001–0.02 %
Γ <sub>19</sub> <i>n</i> γ, helicity=3/2	0.002–0.05 %

*N*(2060) BRANCHING RATIOS

Γ( <i>N</i> π)/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT	Γ <sub>1</sub> /Γ
VALUE (%)				
11 ± 2	SOKHOYAN	15A DPWA	Multichannel	
10 ± 3	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$	
7 ± 2	HOEHLER	79 IPWA	$\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8 ± 2	ANISOVICH	12A DPWA	Multichannel	
9 ± 2	SHRESTHA	12A DPWA	Multichannel	
13 ± 4	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$	

Γ( <i>N</i> η)/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT	Γ <sub>2</sub> /Γ
VALUE (%)				
4 ± 2	ANISOVICH	12A DPWA	Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1	SHRESTHA	12A DPWA	Multichannel	
0.2 ± 1.0	BATINIC	10 DPWA	$\pi N \rightarrow N\pi, N\eta$	

Γ(Σ <i>K</i> )/Γ <sub>total</sub>	DOCUMENT ID	TECN	COMMENT	Γ <sub>4</sub> /Γ
VALUE (%)				
3 ± 2	ANISOVICH	12A DPWA	Multichannel	

See key on page 601

# Baryon Particle Listings

## $N(2060)$ , $N(2100)$

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE (%)				
7 ± 3	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
40 ± 13	SHRESTHA	12A	DPWA Multichannel	

$\Gamma(N\rho, S=1/2, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
VALUE (%)				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
21 ± 15	SHRESTHA	12A	DPWA Multichannel	

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
VALUE (%)				
6 ± 3	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{11}/\Gamma$
VALUE (%)				
9 ± 5	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1520)\pi, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{12}/\Gamma$
VALUE (%)				
15 ± 6	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1680)\pi, S\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{13}/\Gamma$
VALUE (%)				
15 ± 7	SOKHOYAN	15A	DPWA Multichannel	

### $N(2060)$ PHOTON DECAY AMPLITUDES AT THE POLE

$N(2060) \rightarrow \rho\gamma, \text{ helicity-1/2 amplitude } A_{1/2}$	DOCUMENT ID	TECN	COMMENT
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )		
0.064 ± 0.010	12 ± 8	SOKHOYAN	15A DPWA Multichannel

$N(2060) \rightarrow \rho\gamma, \text{ helicity-3/2 amplitude } A_{3/2}$	DOCUMENT ID	TECN	COMMENT
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )		
0.060 ± 0.020	13 ± 10	SOKHOYAN	15A DPWA Multichannel

### $N(2060)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

$N(2060) \rightarrow \rho\gamma, \text{ helicity-1/2 amplitude } A_{1/2}$	DOCUMENT ID	TECN	COMMENT
VALUE ( $\text{GeV}^{-1/2}$ )			
0.062 ± 0.010	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.018 ± 0.004	SHRESTHA	12A	DPWA Multichannel

$N(2060) \rightarrow \rho\gamma, \text{ helicity-3/2 amplitude } A_{3/2}$	DOCUMENT ID	TECN	COMMENT
VALUE ( $\text{GeV}^{-1/2}$ )			
0.062 ± 0.020	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.010 ± 0.004	SHRESTHA	12A	DPWA Multichannel

$N(2060) \rightarrow n\gamma, \text{ helicity-1/2 amplitude } A_{1/2}$	DOCUMENT ID	TECN	COMMENT
VALUE ( $\text{GeV}^{-1/2}$ )			
0.025 ± 0.011	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.012 ± 0.017	SHRESTHA	12A	DPWA Multichannel

$N(2060) \rightarrow n\gamma, \text{ helicity-3/2 amplitude } A_{3/2}$	DOCUMENT ID	TECN	COMMENT
VALUE ( $\text{GeV}^{-1/2}$ )			
−0.037 ± 0.017	ANISOVICH	13B	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.023 ± 0.023	SHRESTHA	12A	DPWA Multichannel

### $N(2060)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $N(2060)$ REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan et al.	(CBELSA/TAPS Collab.)
SVARC	14	PR C89 045205	A. Svarc et al.	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich et al.	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich et al.	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
BATINIC	10	PR C82 038203	M. Batinic et al.	(ZAGR)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky et al.	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky et al.	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hoehler et al.	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$N(2100) 1/2^+$	$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$ Status: *
OMITTED FROM SUMMARY TABLE	

### $N(2100)$ POLE POSITION

REAL PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
2120 ± 25	SOKHOYAN	15A	DPWA Multichannel
2052 ± 6 ± 3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2120 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2120 ± 47	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
1810	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
290 ± 30	SOKHOYAN	15A	DPWA Multichannel
337 ± 10 ± 4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
240 ± 80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
346 ± 80	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
622	VRANA	00	DPWA Multichannel

### $N(2100)$ ELASTIC POLE RESIDUE

MODULUS $ r $	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
23 ± 5	SOKHOYAN	15A	DPWA Multichannel
30 ± 1 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
14 ± 7	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
33	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

PHASE $\theta$	DOCUMENT ID	TECN	COMMENT
VALUE ( $^\circ$ )			
−70 ± 25	SOKHOYAN	15A	DPWA Multichannel
−92 ± 3 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
35 ± 25	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−59	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

### $N(2100)$ INELASTIC POLE RESIDUE

Normalized residue in $N\pi \rightarrow N(2100) \rightarrow \Delta(1232)\pi$	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE ( $^\circ$ )		
0.11 ± 0.05	20 ± 60	SOKHOYAN	15A DPWA Multichannel

Normalized residue in $N\pi \rightarrow N(2100) \rightarrow N\sigma$	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE ( $^\circ$ )		
0.18 ± 0.06	125 ± 25	SOKHOYAN	15A DPWA Multichannel

Normalized residue in $N\pi \rightarrow N(2100) \rightarrow N(1535)\pi$	DOCUMENT ID	TECN	COMMENT
MODULUS	PHASE ( $^\circ$ )		
0.22 ± 0.06	−40 ± 25	SOKHOYAN	15A DPWA Multichannel

### $N(2100)$ BREIT-WIGNER MASS

$\approx 2100$ OUR ESTIMATE	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
2115 ± 20	SOKHOYAN	15A	DPWA Multichannel
2125 ± 75	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2050 ± 20	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2157 ± 42	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
2068 ± 3 <sup>+15</sup> <sub>−40</sub>	ABLIKIM	06K	BES2 $J/\psi \rightarrow (p\pi^-)\bar{n}$
2084 ± 93	VRANA	00	DPWA Multichannel

### $N(2100)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
290 ± 20	SOKHOYAN	15A	DPWA Multichannel
260 ± 100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
200 ± 30	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
355 ± 88	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
165 ± 14 ± 40	ABLIKIM	06K	BES2 $J/\psi \rightarrow (p\pi^-)\bar{n}$
1077 ± 643	VRANA	00	DPWA Multichannel



See key on page 601

Baryon Particle Listings  
*N*(2120), *N*(2190)

2120±35	GUTZ	14	DPWA	Multichannel
2150±60	ANISOVICH	12A	DPWA	Multichannel

*N*(2120) BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
340±35	SOKHOYAN	15A	DPWA Multichannel
300±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (higher <i>m</i> )
265±40	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
340±35	GUTZ	14	DPWA Multichannel
330±45	ANISOVICH	12A	DPWA Multichannel

*N*(2120) DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ <i>N</i> π	5–15 %
$\Gamma_2$ <i>N</i> ππ	50–95 %
$\Gamma_3$ Δ(1232)π	40–90 %
$\Gamma_4$ Δ(1232)π, <i>S</i> -wave	30–70 %
$\Gamma_5$ Δ(1232)π, <i>D</i> -wave	8–32 %
$\Gamma_6$ <i>N</i> σ	7–15 %
$\Gamma_7$ <i>N</i> (1535)π	7–23 %
$\Gamma_8$ <i>p</i> γ	0.16–2.1 %
$\Gamma_9$ <i>p</i> γ, helicity=1/2	0.07–0.80 %
$\Gamma_{10}$ <i>p</i> γ, helicity=3/2	0.09–1.3 %
$\Gamma_{11}$ <i>n</i> γ	0.04–0.72 %
$\Gamma_{12}$ <i>n</i> γ, helicity=1/2	0.04–0.60 %
$\Gamma_{13}$ <i>n</i> γ, helicity=3/2	0.001–0.12 %

*N*(2120) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
5±3	SOKHOYAN	15A	DPWA Multichannel	
14±7	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$ (higher <i>m</i> )	
6±2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
5±3	GUTZ	14	DPWA Multichannel	
6±2	ANISOVICH	12A	DPWA Multichannel	
$\Gamma(\Delta(1232)\pi, S\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
VALUE (%)				
50±20	SOKHOYAN	15A	DPWA Multichannel	
$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE (%)				
20±12	SOKHOYAN	15A	DPWA Multichannel	
$\Gamma(N\sigma)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
VALUE (%)				
11±4	SOKHOYAN	15A	DPWA Multichannel	
$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE (%)				
15±8	GUTZ	14	DPWA Multichannel	

*N*(2120) PHOTON DECAY AMPLITUDES AT THE POLE

<i>N</i> (2120) → <i>p</i> γ, helicity-1/2 amplitude <i>A</i> <sub>1/2</sub>				
MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.130±0.045	−40±25	SOKHOYAN	15A	DPWA Multichannel
<i>N</i> (2120) → <i>p</i> γ, helicity-3/2 amplitude <i>A</i> <sub>3/2</sub>				
MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.160±0.060	−30±15	SOKHOYAN	15A	DPWA Multichannel

*N*(2120) BREIT-WIGNER PHOTON DECAY AMPLITUDES

<i>N</i> (2120) → <i>p</i> γ, helicity-1/2 amplitude <i>A</i> <sub>1/2</sub>				
<i>VALUE</i> (GeV <sup>−1/2</sup> )	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
0.130±0.050	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.130±0.050	GUTZ	14	DPWA	Multichannel
<i>N</i> (2120) → <i>p</i> γ, helicity-3/2 amplitude <i>A</i> <sub>3/2</sub>				
<i>VALUE</i> (GeV <sup>−1/2</sup> )	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
0.160±0.065	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.160±0.065	GUTZ	14	DPWA	Multichannel

*N*(2120) → *n*γ, helicity-1/2 amplitude *A*<sub>1/2</sub>

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.110±0.045	ANISOVICH	13B	DPWA Multichannel

*N*(2120) → *n*γ, helicity-3/2 amplitude *A*<sub>3/2</sub>

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.040±0.030	ANISOVICH	13B	DPWA Multichannel

*N*(2120) REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT)

*N*(2190) 7/2<sup>−</sup>

$I(J^P) = \frac{1}{2}(\frac{7}{2}^-)$  Status: \* \* \* \*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

*N*(2190) POLE POSITION

REAL PART	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
<b>2050 to 2100 (≈ 2075) OUR ESTIMATE</b>			
2150±25	SOKHOYAN	15A	DPWA Multichannel
2079±4±9	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2070	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2042	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
2100±50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2150±25	ANISOVICH	12A	DPWA Multichannel
2062	SHRESTHA	12A	DPWA Multichannel
2063±32	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
2107	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>400 to 520 (≈ 450) OUR ESTIMATE</b>			
325±25	SOKHOYAN	15A	DPWA Multichannel
509±7±16	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
520	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
482	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
400±160	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
330±30	ANISOVICH	12A	DPWA Multichannel
428	SHRESTHA	12A	DPWA Multichannel
330±101	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
380	VRANA	00	DPWA Multichannel

*N*(2190) ELASTIC POLE RESIDUE

MODULUS   <i>r</i>	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)			
<b>25 to 70 (≈ 50) OUR ESTIMATE</b>			
30±4	SOKHOYAN	15A	DPWA Multichannel
54±1±3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
72	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
45	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
25±10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
30±5	ANISOVICH	12A	DPWA Multichannel
34	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

PHASE θ	DOCUMENT ID	TECN	COMMENT
VALUE (°)			
<b>−30 to 30 (≈ 0) OUR ESTIMATE</b>			
28±10	SOKHOYAN	15A	DPWA Multichannel
−18±1±3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−32	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−30±50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
30±10	ANISOVICH	12A	DPWA Multichannel
−19	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$

*N*(2190) INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{\text{pole}}/2$ .

Normalized residue in <i>N</i> π → <i>N</i> (2190) → Λ <i>K</i>				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.03±0.01	20±15	ANISOVICH	12A	DPWA Multichannel

# Baryon Particle Listings

## $N(2190)$

### Normalized residue in $N\pi \rightarrow N(2190) \rightarrow \Delta(1232)\pi$ , $D$ -wave

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
0.27±0.04	−165 ± 20	SOKHOYAN	15A	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow N(2190) \rightarrow N\sigma$

MODULUS	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
0.13±0.05	50 ± 15	SOKHOYAN	15A	DPWA Multichannel

### $N(2190)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2100 to 2200 (<math>\approx 2190</math>) OUR ESTIMATE</b>			
2205 ± 18	SOKHOYAN	15A	DPWA Multichannel
2152.4 ± 1.4	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2200 ± 70	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2140 ± 12	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2180 ± 20	ANISOVICH	12A	DPWA Multichannel
2150 ± 26	SHRESTHA	12A	DPWA Multichannel
2125 ± 61	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
2168 ± 18	VRANA	00	DPWA Multichannel

### $N(2190)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>300 to 700 (<math>\approx 500</math>) OUR ESTIMATE</b>			
355 ± 30	SOKHOYAN	15A	DPWA Multichannel
484 ± 13	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
500 ± 150	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
390 ± 30	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
335 ± 40	ANISOVICH	12A	DPWA Multichannel
500 ± 74	SHRESTHA	12A	DPWA Multichannel
381 ± 160	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
453 ± 101	VRANA	00	DPWA Multichannel

### $N(2190)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	10–20 %
$\Gamma_2$ $N\eta$	seen
$\Gamma_3$ $\Lambda K$	0.2–0.8;%
$\Gamma_4$ $N\pi\pi$	22–80;%
$\Gamma_5$ $\Delta(1232)\pi$	
$\Gamma_6$ $\Delta(1232)\pi$ , $D$ -wave	19–31 %
$\Gamma_7$ $N\rho$	
$\Gamma_8$ $N\rho$ , $S=3/2$ , $D$ -wave	seen
$\Gamma_9$ $N\sigma$	3–9 %
$\Gamma_{10}$ $p\gamma$	0.014–0.077 %
$\Gamma_{11}$ $p\gamma$ , helicity=1/2	0.013–0.062;%
$\Gamma_{12}$ $p\gamma$ , helicity=3/2	0.001–0.014;%
$\Gamma_{13}$ $n\gamma$	<0.04 %
$\Gamma_{14}$ $n\gamma$ , helicity=1/2	<0.01;%
$\Gamma_{15}$ $n\gamma$ , helicity=3/2	<0.03 %

### $N(2190)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
10 to 20 OUR ESTIMATE			
16 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel
23.8 $\pm$ 0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
12 $\pm$ 6	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
14 $\pm$ 2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
16 $\pm$ 2	ANISOVICH	12A	DPWA Multichannel
20 $\pm$ 1	SHRESTHA	12A	DPWA Multichannel
18 $\pm$ 12	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
20 $\pm$ 4	VRANA	00	DPWA Multichannel
$\Gamma(N\eta)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2 $\pm$ 1	SHRESTHA	12A	DPWA Multichannel
0.1 $\pm$ 0.3	BATINIC	10	DPWA $\pi N \rightarrow N\pi, N\eta$
0 $\pm$ 1	VRANA	00	DPWA Multichannel

$\Gamma(\Lambda K)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$0.5 \pm 0.3$	ANISOVICH	12A	DPWA Multichannel
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
<1	SHRESTHA	12A	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$25 \pm 6$	SOKHOYAN	15A	DPWA Multichannel

$\Gamma(N\rho, S=3/2, D\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
29±28	VRANA	00	DPWA	Multichannel

$\Gamma(N\sigma)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
6±3	SOKHOYAN	15A	DPWA	Multichannel

### $N(2190)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $N(2190) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
0.068 ± 0.005	−170 ± 12	SOKHOYAN	15A	DPWA Multichannel

#### $N(2190) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$

MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^{\circ}$ )	DOCUMENT ID	TECN	COMMENT
0.025 ± 0.010	22 ± 10	SOKHOYAN	15A	DPWA Multichannel

### $N(2190)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $N(2190) \rightarrow p\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
−0.071 ± 0.006	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.065 ± 0.008	ANISOVICH	12A	DPWA Multichannel

#### $N(2190) \rightarrow p\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
0.027 ± 0.010	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.035 ± 0.017	ANISOVICH	12A	DPWA Multichannel

#### $N(2190) \rightarrow p\gamma$ , ratio of helicity amplitudes $A_{3/2}/A_{1/2}$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.17 ± 0.15	WILLIAMS	09	IPWA $\gamma p \rightarrow p\omega$

#### $N(2190) \rightarrow n\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
−0.015 ± 0.013	ANISOVICH	13B	DPWA Multichannel

#### $N(2190) \rightarrow n\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
−0.034 ± 0.022	ANISOVICH	13B	DPWA Multichannel

### $N(2190)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $N(2190)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	13B	EPJ A49 67	A.V. Anisovich <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
BATINIC	10	PR C82 038203	M. Batinic <i>et al.</i>	(ZAGR)
WILLIAMS	09	PR C80 065209	M. Williams <i>et al.</i>	(JLab CLAS Collab.)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	79	$\pi N$ Newsletter 9 1	G. Hohlner	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
HOEHLER	79	PDAT 12-1	G. Hohlner <i>et al.</i>	(KARLT) IUP
Also		Toronto Conf. 3	R. Koch	(KARLT) IUP

$$N(2220) \ 9/2^+$$

$$I(J^P) = \frac{1}{2}(\frac{9}{2}^+) \text{ Status: } ****$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### N(2220) POLE POSITION

## REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2130 to 2200 (<math>\approx 2170</math>) OUR ESTIMATE</b>			
2127 $\pm$ 3 $\pm$ 24	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2150 $\pm$ 35	ANISOVICH	12A	DPWA Multichannel
1999	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2135	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
2160 $\pm$ 80	CUT KOSKY	80	IPWA $\pi N \rightarrow \pi N$

**-2x|IMAGINARY PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>400 to 560 (<math>\approx 480</math>) OUR ESTIMATE</b>			
$380 \pm 7 \pm 22$	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
$440 \pm 40$	ANISOVICH	12A	DPWA Multichannel
372	ARNDT	06	DPWA $\pi \rightarrow \pi N, \eta N$
400	HOEHLER	93	ARGD $\pi \rightarrow \pi N$
$480 \pm 100$	CUT KOSKY	80	IPWA $\pi \rightarrow \pi N$

### ***N*(2220) ELASTIC POLE RESIDUE**

**MODULUS  $|r|$**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>35 to 60 (<math>\approx 45</math>) OUR ESTIMATE</b>			
$38 \pm 1 \pm 5$	<sup>1</sup> SVARC	14	$\pi N \rightarrow \pi N$
$60 \pm 12$	ANISOVICH	12A	DPWA Multichannel
33	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
40	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
$45 \pm 20$	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

## PHASE 0

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>-35 to -60 (<math>\approx -50</math>) OUR ESTIMATE</b>			
-52 $\pm$ 1 $\pm$ 14	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
-58 $\pm$ 12	ANISOVICH	12A	DPWA Multichannel
-33	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
-50	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
-45 $\pm$ 25	CUT KOSKY	80	IPWA $\pi N \rightarrow \pi N$

### $N(2220)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2200 to 2300 (<math>\approx</math> 2250) OUR ESTIMATE</b>			
2316.3 $\pm$ 2.9	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2230 $\pm$ 80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2205 $\pm$ 10	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

### ***N*(2220) BREIT-WIGNER WIDTH**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>350 to 500 (<math>\approx 400</math>) OUR ESTIMATE</b>			
633 $\pm$ 17	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
500 $\pm$ 150	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
365 $\pm$ 30	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

## N(2220) DECAY MODES

The following branching fractions are our estimates, not fits or averages.

	Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$N\pi$	15–25 %

### N(2220) BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>15 to 25 OUR ESTIMATE</b>				
24 $\pm$ 5	ANISOVICH	12A	DPWA	Multichannel
24.6 $\pm$ 0.1	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
15 $\pm$ 3	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
18.0 $\pm$ 1.5	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$

## N(2220) FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

## N(2220) REFERENCES

For early references, see *Physics Letters* **111B** 1 (1982).

PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svanc <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
ARNDT	62A	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
HOEHLER	93	$\pi$ N Newsletter 9	G. Hoehler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) UP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) UP
HOEHLER	79	PDAT 12-1	G. Hoehler <i>et al.</i>	(KARL) UP
Also		Toronto Conf. 3	R. Koch	(KARLT) UP

$$N(2250) \ 9/2^-$$

$$I(J^P) = \frac{1}{2}(\frac{9}{2}-) \text{ Status: } ****$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### N(2250) POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2150 to 2250 (<math>\approx 2200</math>) OUR ESTIMATE</b>			
2157 $\pm$ 3 $\pm$ 14	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2195 $\pm$ 45	ANISOVICH	12A	DPWA Multichannel
2217	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2187	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
2150 $\pm$ 50	CUTKOSKY	80	JPWA $\pi N \rightarrow \pi N$

**-2x|IMAGINARY PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>350 to 550 (<math>\approx 450</math>) OUR ESTIMATE</b>			
412 $\pm$ 7 $\pm$ 44	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
470 $\pm$ 50	ANISOVICH	12A DPWA	Multichannel
431	ARNDT	06 DPWA	$\pi N \rightarrow \pi N, \eta N$
388	HOEHLER	93 SPED	$\pi N \rightarrow \pi N$
360 $\pm$ 100	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$

### ***N*(2250) ELASTIC POLE RESIDUE**

**MODULUS  $|r|$**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>20 to 30 (<math>\approx 25</math>) OUR ESTIMATE</b>			
$24 \pm 1 \pm 5$	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
$26 \pm 5$	ANISOVICH	12A DPWA	Multichannel
21	ARNDT	06 DPWA	$\pi N \rightarrow \pi N, \eta N$
21	HOEHLER	93 SPED	$\pi N \rightarrow \pi N$
$20 \pm 6$	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$

### PHASE 0

VALUE (°)	DOCUMENT ID	TECN	COMMENT
-20 to -60 ( $\approx -40$ ) OUR ESTIMATE			
-62 $\pm$ 1 $\pm$ 11	<sup>1</sup> SVARC	14 L+P	$\pi N \rightarrow \pi N$
-38 $\pm$ 25	ANISOVICH	12A DPWA	Multichannel
-20	ARNDT	06 DPWA	$\pi N \rightarrow \pi N, \eta N$
-50 $\pm$ 20	CUTKOSKY	80 IPWA	$\pi N \rightarrow \pi N$

### $N(2250)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2250 to 2320 (<math>\approx 2280</math>) OUR ESTIMATE</b>			
2280 $\pm$ 40	ANISOVICH	12A	DPWA Multichannel
2302 $\pm$ 6	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2250 $\pm$ 80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2268 $\pm$ 15	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

### **$N(2250)$ BREIT-WIGNER WIDTH**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>300 to 600 (<math>\approx 500</math>) OUR ESTIMATE</b>			
520 $\pm$ 50	ANISOVICH	12A	DPWA Multichannel
628 $\pm$ 28	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
480 $\pm$ 120	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
300 $\pm$ 40	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

### *N*(2250) DECAY MODES

The following branching fractions are our estimates, not fits or averages.

	Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$N\pi$	5-15 %





N(∼ 3000) BREIT-WIGNER WIDTH				N(∼ 3000) BRANCHING RATIOS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	$\Gamma(N\pi)/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma$
1300 ± 200	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $L_{1,15}$ wave	VALUE (%)	DOCUMENT ID	TECN	COMMENT
1600 ± 200	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $M_{1,17}$ wave	6 ± 2	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $L_{1,15}$ wave
1900 ± 300	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $N_{1,19}$ wave	4.0 ± 1.5	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $M_{1,17}$ wave
				3.0 ± 1.5	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $N_{1,19}$ wave
N(∼ 3000) DECAY MODES				N(∼ 3000) REFERENCES			
Mode				KOCH	80	Toronto Conf. 3	R. Koch (KARLT) IJP
$\Gamma_1$	$N\pi$			HENDRY	78	PRL 41 222	A.W. Hendry (IND, LBL) IJP
				Also		ANP 136 1	A.W. Hendry (IND) IJP

Baryon Particle Listings

$\Delta(1232)$

$\Delta$  BARYONS  
( $S = 0, I = 3/2$ )

$\Delta^{++} = uuu, \Delta^+ = uud, \Delta^0 = udd, \Delta^- = ddd$

$\Delta(1232) \ 3/2^+$

$I(J^P) = \frac{3}{2}(\frac{3}{2}^+)$  Status: \*\*\*\*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

$\Delta(1232)$  POLE POSITIONS

REAL PART, MIXED CHARGES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1209 to 1211 (<math>\approx 1210</math>) OUR ESTIMATE</b>			
1211 $\pm 1 \pm 1$	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
1210.5 $\pm 1.0$	ANISOVICH 12A	DPWA	Multichannel
1211	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
1209	<sup>2</sup> HOEHLER 93	ARGD	$\pi N \rightarrow \pi N$
1210 $\pm 1$	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1212	SHRESTHA 12A	DPWA	Multichannel
1211 $\pm 1$	ANISOVICH 10	DPWA	Multichannel
1210	ARNDT 04	DPWA	$\pi N \rightarrow \pi N, \eta N$
1217	VRANA 00	DPWA	Multichannel
1211	ARNDT 95	DPWA	$\pi N \rightarrow N\pi$
1210	ARNDT 91	DPWA	$\pi N \rightarrow \pi N$ Soln SM90

−2×IMAGINARY PART, MIXED CHARGES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>98 to 102 (<math>\approx 100</math>) OUR ESTIMATE</b>			
98 $\pm 2 \pm 1$	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
99 $\pm 2$	ANISOVICH 12A	DPWA	Multichannel
99	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
100	<sup>2</sup> HOEHLER 93	ARGD	$\pi N \rightarrow \pi N$
100 $\pm 2$	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
98	SHRESTHA 12A	DPWA	Multichannel
100 $\pm 2$	ANISOVICH 10	DPWA	Multichannel
100	ARNDT 04	DPWA	$\pi N \rightarrow \pi N, \eta N$
96	VRANA 00	DPWA	Multichannel
100	ARNDT 95	DPWA	$\pi N \rightarrow N\pi$
100	ARNDT 91	DPWA	$\pi N \rightarrow \pi N$ Soln SM90

REAL PART,  $\Delta(1232)^{++}$

VALUE (MeV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1212.50 $\pm 0.24$	BERNICA 96	Fit to PEDRONI 78

−2×IMAGINARY PART,  $\Delta(1232)^{++}$

VALUE (MeV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
97.37 $\pm 0.42$	BERNICA 96	Fit to PEDRONI 78

REAL PART,  $\Delta(1232)^+$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1211 $\pm 1$ to 1212 $\pm 1$	HANSTEIN 96	DPWA	$\gamma N \rightarrow \pi N$
1206.9 $\pm 0.9$ to 1210.5 $\pm 1.8$	MIROSHNIC... 79		Fit photoproduction

−2×IMAGINARY PART,  $\Delta(1232)^+$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
102 $\pm 2$ to 99 $\pm 2$	<sup>3</sup> HANSTEIN 96	DPWA	$\gamma N \rightarrow \pi N$
111.2 $\pm 2.0$ to 116.6 $\pm 2.2$	MIROSHNIC... 79		Fit photoproduction

REAL PART,  $\Delta(1232)^0$

VALUE (MeV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1213.20 $\pm 0.66$	BERNICA 96	Fit to PEDRONI 78

−2×IMAGINARY PART,  $\Delta(1232)^0$

VALUE (MeV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
104.10 $\pm 1.01$	BERNICA 96	Fit to PEDRONI 78

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.  
<sup>2</sup> See HOEHLER 93 for a detailed discussion of the evidence for and the pole parameters of  $N$  and  $\Delta$  resonances as determined from Argand diagrams of  $\pi N$  elastic partial-wave amplitudes and from plots of the speeds with which the amplitudes traverse the diagrams.  
<sup>3</sup> The second (lower) value of HANSTEIN 96 here goes with the second (higher) value of the real part in the preceding data block.

$\Delta(1232)$  ELASTIC POLE RESIDUES

ABSOLUTE VALUE, MIXED CHARGES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
50 $\pm 1 \pm 1$	<sup>4</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
51.6 $\pm 0.6$	ANISOVICH 12A	DPWA	Multichannel
52	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
50	HOEHLER 93	ARGD	$\pi N \rightarrow \pi N$
53 $\pm 2$	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
53	ARNDT 04	DPWA	$\pi N \rightarrow \pi N, \eta N$
38	<sup>5</sup> ARNDT 95	DPWA	$\pi N \rightarrow N\pi$
52	ARNDT 91	DPWA	$\pi N \rightarrow \pi N$ Soln SM90

PHASE, MIXED CHARGES

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−46 $\pm 1 \pm 1$	<sup>4</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
−46 $\pm 1$	ANISOVICH 12A	DPWA	Multichannel
−47	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
−48	HOEHLER 93	ARGD	$\pi N \rightarrow \pi N$
−47 $\pm 1$	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−47	ARNDT 04	DPWA	$\pi N \rightarrow \pi N, \eta N$
−22	<sup>5</sup> ARNDT 95	DPWA	$\pi N \rightarrow N\pi$
−31	ARNDT 91	DPWA	$\pi N \rightarrow \pi N$ Soln SM90

<sup>4</sup> Fit to the amplitudes of HOEHLER 79.  
<sup>5</sup> This ARNDT 95 value is in error, as pointed out by HOHLER 01. The corrected value is in line with the ARNDT 91 value (R.A. Arndt, private communication).

$\Delta(1232)$  BREIT-WIGNER MASSES

MIXED CHARGES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1230 to 1234 (<math>\approx 1232</math>) OUR ESTIMATE</b>			
1228 $\pm 2$	ANISOVICH 12A	DPWA	Multichannel
1233.4 $\pm 0.4$	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
1232 $\pm 3$	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
1233 $\pm 2$	HOEHLER 79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1231.1 $\pm 0.2$	SHRESTHA 12A	DPWA	Multichannel
1230 $\pm 2$	ANISOVICH 10	DPWA	Multichannel
1232.9 $\pm 1.2$	ARNDT 04	DPWA	$\pi N \rightarrow \pi N, \eta N$
1228 $\pm 1$	PENNER 02C	DPWA	Multichannel
1234 $\pm 5$	VRANA 00	DPWA	Multichannel
1233	ARNDT 95	DPWA	$\pi N \rightarrow N\pi$
1231 $\pm 1$	MANLEY 92	IPWA	$\pi N \rightarrow \pi N$ & $N\pi\pi$

$\Delta(1232)^{++}$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1230.55 $\pm 0.20$	GRIDNEV 06	DPWA	$\pi N \rightarrow \pi N$
1231.88 $\pm 0.29$	BERNICA 96		Fit to PEDRONI 78
1230.5 $\pm 0.2$	ABAEV 95	IPWA	$\pi N \rightarrow \pi N$
1230.9 $\pm 0.3$	KOCH 80B	IPWA	$\pi N \rightarrow \pi N$
1231.1 $\pm 0.2$	PEDRONI 78		$\pi N \rightarrow \pi N$ 70–370 MeV

$\Delta(1232)^+$  MASS

VALUE (MeV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
1234.9 $\pm 1.4$	MIROSHNIC... 79	Fit photoproduction

$\Delta(1232)^0$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1231.3 $\pm 0.6$	BREITSCHOP...06	CNTR	Using new CHEX data
1233.40 $\pm 0.22$	GRIDNEV 06	DPWA	$\pi N \rightarrow \pi N$
1234.35 $\pm 0.75$	BERNICA 96		Fit to PEDRONI 78
1233.1 $\pm 0.3$	ABAEV 95	IPWA	$\pi N \rightarrow \pi N$
1233.6 $\pm 0.5$	KOCH 80B	IPWA	$\pi N \rightarrow \pi N$
1233.8 $\pm 0.2$	PEDRONI 78		$\pi N \rightarrow \pi N$ 70–370 MeV

$m_{\Delta^0} - m_{\Delta^{++}}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.86 $\pm 0.30$	GRIDNEV 06	DPWA	$\pi N \rightarrow \pi N$
2.25 $\pm 0.68$	BERNICA 96		Fit to PEDRONI 78
2.6 $\pm 0.4$	ABAEV 95	IPWA	$\pi N \rightarrow \pi N$
2.7 $\pm 0.3$	<sup>6</sup> PEDRONI 78		See the masses

<sup>6</sup> Using  $\pi^\pm d$  as well, PEDRONI 78 determine  $(M^- - M^{++}) + (M^0 - M^+)/3 = 4.6 \pm 0.2$  MeV.

See key on page 601

Baryon Particle Listings  
 $\Delta(1232)$

$\Delta(1232)$  BREIT-WIGNER WIDTHS

MIXED CHARGES

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>114 to 120 (<math>\approx 117</math>) OUR ESTIMATE</b>			
110 $\pm$ 3	ANISOVICH	12A	DPWA Multichannel
118.7 $\pm$ 0.6	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
120 $\pm$ 5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
116 $\pm$ 5	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
113.0 $\pm$ 0.5	SHRESTHA	12A	DPWA Multichannel
112 $\pm$ 4	ANISOVICH	10	DPWA Multichannel
118.0 $\pm$ 2.2	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$
106 $\pm$ 1	PENNER	02c	DPWA Multichannel
112 $\pm$ 18	VRANA	00	DPWA Multichannel
114	ARNDT	95	DPWA $\pi N \rightarrow N\pi$
118 $\pm$ 4	MANLEY	92	IPWA $\pi N \rightarrow \pi N$ & $N\pi\pi$

$\Delta(1232)^{++}$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
112.2 $\pm$ 0.7	GRIDNEV	06	DPWA $\pi N \rightarrow \pi N$
109.07 $\pm$ 0.48	BERNICHIA	96	Fit to PEDRONI 78
111.0 $\pm$ 1.0	KOCH	80b	IPWA $\pi N \rightarrow \pi N$
111.3 $\pm$ 0.5	PEDRONI	78	$\pi N \rightarrow \pi N$ 70–370 MeV

$\Delta(1232)^+$  WIDTH

VALUE (MeV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
131.1 $\pm$ 2.4	MIROSHNIC... 79	Fit photoproduction

$\Delta(1232)^0$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
112.5 $\pm$ 1.9	BREITSCHOP..06	CNTR	Using new CHEX data
116.9 $\pm$ 0.7	GRIDNEV	06	DPWA $\pi N \rightarrow \pi N$
117.58 $\pm$ 1.16	BERNICHIA	96	Fit to PEDRONI 78
113.0 $\pm$ 1.5	KOCH	80b	IPWA $\pi N \rightarrow \pi N$
117.9 $\pm$ 0.9	PEDRONI	78	$\pi N \rightarrow \pi N$ 70–370 MeV

$\Delta^0$ - $\Delta^{++}$  WIDTH DIFFERENCE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4.66 $\pm$ 1.0	GRIDNEV	06	DPWA $\pi N \rightarrow \pi N$
8.45 $\pm$ 1.11	BERNICHIA	96	Fit to PEDRONI 78
5.1 $\pm$ 1.0	ABAEV	95	IPWA $\pi N \rightarrow \pi N$
6.6 $\pm$ 1.0	PEDRONI	78	See the widths

$\Delta(1232)$  DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	99.4 %
$\Gamma_2$ $N\gamma$	0.55–0.65 %
$\Gamma_3$ $N\gamma$ , helicity=1/2	0.11–0.13 %
$\Gamma_4$ $N\gamma$ , helicity=3/2	0.44–0.52 %

$\Delta(1232)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>0.994 OUR ESTIMATE</b>				
1.00	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$	
1.0	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
1.0	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.994	SHRESTHA	12A	DPWA Multichannel	
1.0	ANISOVICH	10	DPWA Multichannel	
1.000	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$	
1.00	PENNER	02c	DPWA Multichannel	
1.00 $\pm$ 0.01	VRANA	00	DPWA Multichannel	
1.0	ARNDT	95	DPWA $\pi N \rightarrow N\pi$	
1.0	MANLEY	92	IPWA $\pi N \rightarrow \pi N$ & $N\pi\pi$	

$\Delta(1232)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES

Papers on  $\gamma N$  amplitudes predating 1981 may be found in our 2006 edition, Journal of Physics **G33** 1 (2006).

$\Delta(1232) \rightarrow N\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>-0.135 <math>\pm</math> 0.006 OUR ESTIMATE</b>			
-0.131 $\pm$ 0.004	ANISOVICH	12A	DPWA Multichannel
-0.139 $\pm$ 0.002	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
-0.139 $\pm$ 0.004	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
-0.137 $\pm$ 0.005	AHRENS	04A	DPWA $\vec{\gamma}\vec{p} \rightarrow N\pi$
-0.1357 $\pm$ 0.0013 $\pm$ 0.0037	BLANPIED	01	LEGS $\gamma p \rightarrow p\gamma, p\pi^0, n\pi^+$
-0.131 $\pm$ 0.001	BECK	00	IPWA $\vec{\gamma}p \rightarrow p\pi^0, n\pi^+$
-0.140 $\pm$ 0.005	KAMALOV	99	DPWA $\gamma N \rightarrow \pi N$
-0.1294 $\pm$ 0.0013	HANSTEIN	98	IPWA $\gamma N \rightarrow \pi N$
-0.1278 $\pm$ 0.0012	DAVIDSON	97	DPWA $\gamma N \rightarrow \pi N$
-0.135 $\pm$ 0.016	DAVIDSON	91B	FIT $\gamma N \rightarrow \pi N$
-0.145 $\pm$ 0.015	CRAWFORD	83	IPWA $\gamma N \rightarrow \pi N$
-0.138 $\pm$ 0.004	AWAJI	81	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.137 $\pm$ 0.001	SHRESTHA	12A	DPWA Multichannel
-0.136 $\pm$ 0.005	ANISOVICH	10	DPWA Multichannel
-0.140	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
-0.129 $\pm$ 0.001	ARNDT	02	DPWA $\gamma p \rightarrow N\pi$
-0.128	PENNER	02D	DPWA Multichannel
-0.1312	HANSTEIN	98	DPWA $\gamma N \rightarrow \pi N$
-0.135 $\pm$ 0.005	ARNDT	97	IPWA $\gamma N \rightarrow \pi N$
-0.141 $\pm$ 0.005	ARNDT	96	IPWA $\gamma N \rightarrow \pi N$
-0.143 $\pm$ 0.004	LI	93	IPWA $\gamma N \rightarrow \pi N$
-0.140 $\pm$ 0.007	DAVIDSON	90	FIT See DAVIDSON 91B

$\Delta(1232) \rightarrow N\gamma$ , helicity-3/2 amplitude  $A_{3/2}$

VALUE (GeV $^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>-0.255 <math>\pm</math> 0.005 OUR ESTIMATE</b>			
-0.254 $\pm$ 0.005	ANISOVICH	12A	DPWA Multichannel
-0.262 $\pm$ 0.003	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
-0.258 $\pm$ 0.005	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
-0.256 $\pm$ 0.003	AHRENS	04A	DPWA $\vec{\gamma}\vec{p} \rightarrow N\pi$
-0.2669 $\pm$ 0.0016 $\pm$ 0.0078	BLANPIED	01	LEGS $\gamma p \rightarrow p\gamma, p\pi^0, n\pi^+$
-0.251 $\pm$ 0.001	BECK	00	IPWA $\vec{\gamma}p \rightarrow p\pi^0, n\pi^+$
-0.258 $\pm$ 0.006	KAMALOV	99	DPWA $\gamma N \rightarrow \pi N$
-0.2466 $\pm$ 0.0013	HANSTEIN	98	IPWA $\gamma N \rightarrow \pi N$
-0.2524 $\pm$ 0.0013	DAVIDSON	97	DPWA $\gamma N \rightarrow \pi N$
-0.251 $\pm$ 0.033	DAVIDSON	91B	FIT $\gamma N \rightarrow \pi N$
-0.263 $\pm$ 0.026	CRAWFORD	83	IPWA $\gamma N \rightarrow \pi N$
-0.259 $\pm$ 0.006	AWAJI	81	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.251 $\pm$ 0.001	SHRESTHA	12A	DPWA Multichannel
-0.267 $\pm$ 0.008	ANISOVICH	10	DPWA Multichannel
-0.265	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
-0.243 $\pm$ 0.001	ARNDT	02	DPWA $\gamma p \rightarrow N\pi$
-0.247	PENNER	02D	DPWA Multichannel
-0.2522	HANSTEIN	98	DPWA $\gamma N \rightarrow \pi N$
-0.250 $\pm$ 0.008	ARNDT	97	IPWA $\gamma N \rightarrow \pi N$
-0.261 $\pm$ 0.005	ARNDT	96	IPWA $\gamma N \rightarrow \pi N$
-0.262 $\pm$ 0.004	LI	93	IPWA $\gamma N \rightarrow \pi N$
-0.254 $\pm$ 0.011	DAVIDSON	90	FIT See DAVIDSON 91B

$\Delta(1232) \rightarrow N\gamma$ ,  $E_2/M_1$  ratio

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.025 <math>\pm</math> 0.005 OUR ESTIMATE</b>			
-0.0274 $\pm$ 0.0003 $\pm$ 0.0030	AHRENS	04A	DPWA $\vec{\gamma}\vec{p} \rightarrow N\pi$
-0.020 $\pm$ 0.002	ARNDT	02	DPWA $\gamma p \rightarrow N\pi$
-0.0307 $\pm$ 0.0026 $\pm$ 0.0024	BLANPIED	01	LEGS $\gamma p \rightarrow p\gamma, p\pi^0, n\pi^+$
-0.016 $\pm$ 0.004 $\pm$ 0.002	GALLER	01	DPWA $\gamma p \rightarrow \gamma p$
-0.025 $\pm$ 0.001 $\pm$ 0.002	BECK	00	IPWA $\vec{\gamma}p \rightarrow p\pi^0, n\pi^+$
-0.0233 $\pm$ 0.0017	HANSTEIN	98	IPWA $\gamma N \rightarrow \pi N$
-0.015 $\pm$ 0.005	7 ARNDT	97	IPWA $\gamma N \rightarrow \pi N$
-0.0319 $\pm$ 0.0024	DAVIDSON	97	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.022	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
-0.026	PENNER	02D	DPWA Multichannel
-0.0254 $\pm$ 0.0010	HANSTEIN	98	DPWA $\gamma N \rightarrow \pi N$
-0.025 $\pm$ 0.002 $\pm$ 0.002	BECK	97	IPWA $\gamma N \rightarrow \pi N$
-0.030 $\pm$ 0.003 $\pm$ 0.002	BLANPIED	97	DPWA $\gamma N \rightarrow \pi N, \gamma N$
-0.027 $\pm$ 0.003 $\pm$ 0.001	KHANDAKER	95	DPWA $\gamma N \rightarrow \pi N$
-0.015 $\pm$ 0.005	WORKMAN	92	IPWA $\gamma N \rightarrow \pi N$
-0.0157 $\pm$ 0.0072	DAVIDSON	91B	FIT $\gamma N \rightarrow \pi N$
-0.0107 $\pm$ 0.0037	DAVIDSON	90	FIT $\gamma N \rightarrow \pi N$
-0.015 $\pm$ 0.002	DAVIDSON	86	FIT $\gamma N \rightarrow \pi N$
+0.037 $\pm$ 0.004	TANABE	85	FIT $\gamma N \rightarrow \pi N$

$\Delta(1232) \rightarrow N\gamma$ , absolute value of  $E_2/M_1$  ratio at pole

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.065 $\pm$ 0.007	ARNDT	97	DPWA $\gamma N \rightarrow \pi N$
0.058	HANSTEIN	96	DPWA $\gamma N \rightarrow \pi N$

# Baryon Particle Listings

## $\Delta(1232)$ , $\Delta(1600)$

### $\Delta(1232) \rightarrow N\gamma$ , phase of $E_2/M_1$ ratio at pole

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−122 ±5	ARNDT 97	DPWA	$\gamma N \rightarrow \pi N$
−127.2	HANSTEIN 96	DPWA	$\gamma N \rightarrow \pi N$

<sup>7</sup> This ARNDT 97 value is very sensitive to the database being fitted. The result is from a fit to the full pion photoproduction database, apart from the BLANPIED 97 cross-section measurements.

### $\Delta(1232)$ MAGNETIC MOMENTS

#### $\Delta(1232)^{++}$ MAGNETIC MOMENT

The values are extracted from UCLA and SIN data on  $\pi^+ p$  bremsstrahlung using a variety of different theoretical approximations and methods. Our estimate is *only* a rough guess of the range we expect the moment to lie within.

VALUE ( $\mu_N$ )	DOCUMENT ID	TECN	COMMENT
<b>3.7 to 7.5 OUR ESTIMATE</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6.14±0.51	LOPEZCAST... 01	DPWA	$\pi^+ p \rightarrow \pi^+ p \gamma$
4.52±0.50±0.45	BOSSHARD 91		$\pi^+ p \rightarrow \pi^+ p \gamma$ (SIN data)
3.7 to 4.2	LIN 91B		$\pi^+ p \rightarrow \pi^+ p \gamma$ (from UCLA data)
4.6 to 4.9	LIN 91B		$\pi^+ p \rightarrow \pi^+ p \gamma$ (from SIN data)
5.6 to 7.5	WITTMAN 88		$\pi^+ p \rightarrow \pi^+ p \gamma$ (from UCLA data)
6.9 to 9.8	HELLER 87		$\pi^+ p \rightarrow \pi^+ p \gamma$ (from UCLA data)
4.7 to 6.7	NEFKENS 78		$\pi^+ p \rightarrow \pi^+ p \gamma$ (UCLA data)

#### $\Delta(1232)^+$ MAGNETIC MOMENT

VALUE ( $\mu_N$ )	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
2.7 <sup>+1.0</sup> <sub>−1.3</sub> ± 1.5 ± 3	<sup>8</sup> KOTULLA 02	$\gamma p \rightarrow p \pi^0 \gamma'$
<sup>8</sup> The second error is systematic, the third is an estimate of theoretical uncertainties.		

### $\Delta(1232)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

PDG 14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC 14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH 12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA 12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN 12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
ANISOVICH 10	EPJ A44 203	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
DRECHSEL 07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER 07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT 06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
BREITSCHOP... 06	PL B639 424	J. Breitschopf <i>et al.</i>	(TUBIN, HEBR, CSUS)
GRIDNEV 06	PAN 69 1542	A.B. Gridnev <i>et al.</i>	(PNPI, BONN, GWU)
PDG 06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
AHRENS 04A	EPJ A21 323	J. Ahrens <i>et al.</i>	(Mainz GDH, A2 Collab.)
ARNDT 04	PR C69 035213	R.A. Arndt <i>et al.</i>	(GWU, TRIU)
ARNDT 02	PR C66 055213	R. A. Arndt <i>et al.</i>	(GWU)
KOTULLA 02	PRL 89 272001	M. Kotulla <i>et al.</i>	(MAMI TAPS Collab.)
PENNER 02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER 02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
BLANPIED 01	PR C64 025203	G. Blanpied <i>et al.</i>	(BNL LEGS Collab.)
GALLER 01	PL B503 245	G. Galler <i>et al.</i>	(Mainz LARA Collab.)
HOHLER 01	NSTAR 2001 185	G. Hohler	(KARL)
LOPEZCAST... 01	NP A697 440	G. Lopez Castro, A. Mariano	
Also	NP A697 440	G. Lopez Castro, A. Mariano	
BECK 00	PR C61 035204	R. Beck <i>et al.</i>	(Mainz Microtron DAPHNE Col.)
VRANA 00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
KAMALOV 99	PRL 83 4494	S.S. Kamalov, S.N. Yang	(Taiwan U.)
HANSTEIN 98	NP A632 561	O. Hanstein, D. Drechsel, L. Tiator	
ARNDT 97	PR C56 577	R.A. Arndt, I.I. Strakovsky, R.L. Workman	(VPI)
BECK 97	PRL 78 606	R. Beck <i>et al.</i>	(MANZ, SACL, PAVI, GLAS)
Also	PRL 79 4510	R.L. Beck, H.P. Krahn	(MANZ)
Also	PRL 79 4512	R.L. Beck, H.P. Krahn	(MANZ)
Also	PRL 79 4515 (erratum)	R.L. Beck <i>et al.</i>	(MANZ, SACL, PAVI, GLAS)
BLANPIED 97	PRL 79 4337	G.S. Blanpied <i>et al.</i>	(LEGS Collab.)
DAVIDSON 97	PRL 79 4509	R.M. Davidson, N.C.A. Mukhopadhyay	(RPI)
ARNDT 96	PR C53 430	R.A. Arndt, I.I. Strakovsky, R.L. Workman	(VPI)
BERNICHIA 96	NP A597 623	A. Bernichia, G. Lopez Castro, J. Pestieau	(LOUV+)
HANSTEIN 96	PL B385 45	O. Hanstein, D. Drechsel, L. Tiator	(MANZ)
ABAEV 95	ZPHY A352 85	V.V. Abaev, S.P. Kruglov	(PNPI)
ARNDT 95	PR C52 2120	R.A. Arndt <i>et al.</i>	(VPI, BRCO)
KHANDAKER 95	PR D51 3966	M. Khandaker, A.M. Sandomi	(BNL, VPI)
HOEHLER 93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
LI 93	PR C47 2759	Z.J. Li <i>et al.</i>	(VPI)
MANLEY 92	PR D45 4002	D.M. Manley, E.M. Saleski	(KSA) IJP
Also	PR D30 904	D.M. Manley <i>et al.</i>	(VPI)
WORKMAN 92	PR C46 1546	R. Workman, R.A. Arndt, Z.J. Li	(VPI)
ARNDT 91	PR D43 2131	R.A. Arndt <i>et al.</i>	(VPI, TELE) IJP
BOSSHARD 91	PR D44 1962	A. Bosshard <i>et al.</i>	(ZURI, LBL, VILL+)
Also	PRL 64 2619	A. Bosshard <i>et al.</i>	(CATH, LAUS, LBL+)
DAVIDSON 91B	PR D43 71	R.M. Davidson, N.C. Mukhopadhyay, R.S. Wittman	(RPI)
LIN 91B	PR C44 1819	D.H. Lin, M.K. Liou, Z.M. Ding	(CUNY, CSOK)
Also	PR C43 R930	D. Lin, M.K. Liou	(CUNY)
DAVIDSON 90	PR D42 20	R.M. Davidson, N.C. Mukhopadhyay	(RPI)
WITTMAN 88	PR C37 2075	R. Wittman	(TRIU)
HELLER 87	PR C35 718	L. Heller <i>et al.</i>	(LANL, MIT, ILL)
DAVIDSON 86	PRL 56 804	R.M. Davidson, N.C. Mukhopadhyay, R. Wittman	(RPI)
TANABE 85	PR C31 1876	H. Tanabe, K. Oita	(KOMAB)
CRAWFORD 83	NP B211 1	R.L. Crawford, W.T. Morton	(GLAS)
AWAJI 81	Bonn Conf. 352	N. Awaji, R. Kajikawa	(NAGO)
Also	NP B197 365	K. Fujii <i>et al.</i>	(NAGO)
CUTKOSKY 80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also	PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
KOCH 80B	NP A336 331	R. Koch, E. Pietarinen	(KARLT) IJP
HOEHLER 79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also	Toronto Conf. 3	R. Koch	(KARLT) IJP
MIROSHNIC... 79	SJNP 29 94	I.I. Miroshnichenko <i>et al.</i>	(KFTI) IJP
Also	Translated from YAF 29 188		
NEFKENS 78	PR D18 3911	B.M.K. Nefkens <i>et al.</i>	(UCLA, CATH) IJP
PEDRONI 78	NP A300 321	E. Pedroni <i>et al.</i>	(SIN, ISNG, KARL+) IJP

## $\Delta(1600) 3/2^+$

$$I(J^P) = \frac{3}{2}(\frac{3}{2}^+) \text{ Status: } ***$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $\Delta(1600)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1460 to 1560 (<math>\approx</math> 1510) OUR ESTIMATE</b>			
1515 ± 20	SOKHOYAN 15A	DPWA	Multichannel
1469 ± 10 ± 5	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
1457	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
1550	HOEHLER 93	SPED	$\pi N \rightarrow \pi N$
1550 ± 40	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1498 ± 25	ANISOVICH 12A	DPWA	Multichannel
1599	SHRESTHA 12A	DPWA	Multichannel
1599	VRANA 00	DPWA	Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>200 to 350 (<math>\approx</math> 275) OUR ESTIMATE</b>			
250 ± 30	SOKHOYAN 15A	DPWA	Multichannel
314 ± 18 ± 8	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
400	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
200 ± 60	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
230 ± 50	ANISOVICH 12A	DPWA	Multichannel
211	SHRESTHA 12A	DPWA	Multichannel
312	VRANA 00	DPWA	Multichannel

### $\Delta(1600)$ ELASTIC POLE RESIDUE

#### MODULUS $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>10 to 40 (<math>\approx</math> 25) OUR ESTIMATE</b>			
13 ± 3	SOKHOYAN 15A	DPWA	Multichannel
38 ± 2 ± 2	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
44	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
17 ± 4	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11 ± 6	ANISOVICH 12A	DPWA	Multichannel
<b>PHASE <math>\theta</math></b>			
<b>150 to 210 (<math>\approx</math> 180) OUR ESTIMATE</b>			
−155 ± 20	SOKHOYAN 15A	DPWA	Multichannel
173 ± 5 ± 5	<sup>1</sup> SVARC 14	L+P	$\pi N \rightarrow \pi N$
+147	ARNDT 06	DPWA	$\pi N \rightarrow \pi N, \eta N$
−150 ± 30	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−160 ± 33	ANISOVICH 12A	DPWA	Multichannel

### $\Delta(1600)$ INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow \Delta(1600) \rightarrow \Delta\pi$ , $P$ -wave

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
15 ± 4	30 ± 35	SOKHOYAN 15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
14 ± 10	154 ± 40	ANISOVICH 12A	DPWA	Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1600) \rightarrow \Delta\pi$ , $F$ -wave

MODULUS (%)	DOCUMENT ID	TECN	COMMENT
1.0 ± 0.5	SOKHOYAN 15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.0 ± 0.5	ANISOVICH 12A	DPWA	Multichannel

### $\Delta(1600)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1500 to 1700 (<math>\approx</math> 1600) OUR ESTIMATE</b>			
1520 ± 20	SOKHOYAN 15A	DPWA	Multichannel
1600 ± 50	CUTKOSKY 80	IPWA	$\pi N \rightarrow \pi N$
1522 ± 13	HOEHLER 79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1510 ± 20	ANISOVICH 12A	DPWA	Multichannel
1626 ± 8	SHRESTHA 12A	DPWA	Multichannel
1667 ± 1	PENNER 02C	DPWA	Multichannel
1687 ± 44	VRANA 00	DPWA	Multichannel

See key on page 601

Baryon Particle Listings  
 $\Delta(1600)$ ,  $\Delta(1620)$

$\Delta(1600)$ BREIT-WIGNER WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>220 to 420 (<math>\approx 320</math>) OUR ESTIMATE</b>				
235 $\pm$ 30	SOKHOYAN	15A	DPWA	Multichannel
300 $\pm$ 100	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
220 $\pm$ 40	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
220 $\pm$ 45	ANISOVICH	12A	DPWA	Multichannel
225 $\pm$ 18	SHRESTHA	12A	DPWA	Multichannel
397 $\pm$ 10	PENNER	02c	DPWA	Multichannel
493 $\pm$ 75	VRANA	00	DPWA	Multichannel

$\Delta(1600)$ DECAY MODES		
The following branching fractions are our estimates, not fits or averages.		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\pi$	10–25 %	
$\Gamma_2$ $N\pi\pi$	75–90 %	
$\Gamma_3$ $\Delta(1232)\pi$	73–83 %	
$\Gamma_4$ $\Delta(1232)\pi$ , $P$ -wave	72–82 %	
$\Gamma_5$ $\Delta(1232)\pi$ , $F$ -wave	<2 %	
$\Gamma_6$ $N(1440)\pi$		
$\Gamma_7$ $N(1440)\pi$ , $P$ -wave	seen	
$\Gamma_8$ $N\gamma$	0.001–0.035 %	
$\Gamma_9$ $N\gamma$ , helicity=1/2	0.0–0.02 %	
$\Gamma_{10}$ $N\gamma$ , helicity=3/2	0.001–0.015 %	

$\Delta(1600)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
<b>10 to 25 OUR ESTIMATE</b>			
14 $\pm$ 4	SOKHOYAN	15A	DPWA Multichannel
18 $\pm$ 4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
21 $\pm$ 6	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
12 $\pm$ 5	ANISOVICH	12A	DPWA Multichannel
8 $\pm$ 2	SHRESTHA	12A	DPWA Multichannel
13 $\pm$ 1	PENNER	02c	DPWA Multichannel
28 $\pm$ 5	VRANA	00	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
77 $\pm$ 5	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
78 $\pm$ 6	ANISOVICH	12A	DPWA Multichannel
70 $\pm$ 3	SHRESTHA	12A	DPWA Multichannel
59 $\pm$ 10	VRANA	00	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
<2	SOKHOYAN	15A	DPWA Multichannel

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
22 $\pm$ 3	SHRESTHA	12A	DPWA Multichannel
13 $\pm$ 4	VRANA	00	DPWA Multichannel

$\Delta(1600)$ PHOTON DECAY AMPLITUDES AT THE POLE				
$\Delta(1600) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
0.053 $\pm$ 0.010	130 $\pm$ 15	SOKHOYAN	15A	DPWA Multichannel
$\Delta(1600) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
0.055 $\pm$ 0.010	152 $\pm$ 15	SOKHOYAN	15A	DPWA Multichannel

$\Delta(1600)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES			
$\Delta(1600) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$			
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>–0.045 <math>\pm</math> 0.015 OUR ESTIMATE</b>			
–0.051 $\pm$ 0.010	SOKHOYAN	15A	DPWA Multichannel
–0.018 $\pm$ 0.015	ARNDT	96	IPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–0.050 $\pm$ 0.009	ANISOVICH	12A	DPWA Multichannel
0.006 $\pm$ 0.005	SHRESTHA	12A	DPWA Multichannel
0.0	PENNER	02d	DPWA Multichannel

$\Delta(1600) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$			
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>–0.035 <math>\pm</math> 0.015 OUR ESTIMATE</b>			
–0.055 $\pm$ 0.010	SOKHOYAN	15A	DPWA Multichannel
–0.025 $\pm$ 0.015	ARNDT	96	IPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–0.040 $\pm$ 0.012	ANISOVICH	12A	DPWA Multichannel
0.052 $\pm$ 0.008	SHRESTHA	12A	DPWA Multichannel
–0.024	PENNER	02d	DPWA Multichannel

$\Delta(1600)$ FOOTNOTES	
<sup>1</sup> Fit to the amplitudes of HOEHLER 79.	

$\Delta(1600)$ REFERENCES	
For early references, see Physics Letters <b>111B 1</b> (1982).	
SOKHOYAN 15A PDG 14 SVARC 14 ANISOVICH 12A SHRESTHA 12A ARNDT 06 PENNER 02C PENNER 02D VRANA 00 ARNDT 96 HOEHLER 93 CUTKOSKY 80 Also HOEHLER 79	EPJ A51 95 CPC 38 070001 PR C89 045205 EPJ A48 15 PR C86 055203 PR C74 045205 PR C66 055211 PR C66 055212 PRPL 328 181 PR C53 430 $\pi N$ Newsletter 9 1 Toronto Conf. 19 PR D20 2839 PDAT 12-1 Toronto Conf. 3
V. Sokhoyan <i>et al.</i> K. Olive <i>et al.</i> A. Svane <i>et al.</i> A.V. Anisovich <i>et al.</i> M. Shrestha, D.M. Manley R.A. Arndt <i>et al.</i> G. Penner, U. Mosel G. Penner, U. Mosel T.P. Vrana, S.A. Dytmann, T.-S.H. Lee R.A. Arndt, I.I. Strakovsky, R.L. Workman G. Hohlner R.E. Cutkosky <i>et al.</i> R.E. Cutkosky <i>et al.</i> G. Hohlner <i>et al.</i> R. Koch	(CBELSA/TAPS Collab.) (PDG Collab.) (BONN, PNPI) (KSU) (GWU) (GIES) (GIES) (PITT, ANL) (VPI) (KARL) (CMU, LBL) IJP (CMU, LBL) IJP (KARLT) IJP (KARLT) IJP

$\Delta(1620) 1/2^-$	$I(J^P) = \frac{3}{2}(\frac{1}{2}^-)$ Status: * * * *
Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C <b>38</b> 070001 (2014).	

$\Delta(1620)$ POLE POSITION			
REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1590 to 1610 (<math>\approx 1600</math>) OUR ESTIMATE</b>			
1597 $\pm$ 5	SOKHOYAN	15A	DPWA Multichannel
1603 $\pm$ 7 $\pm$ 2	SVARC	14	L+P $\pi N \rightarrow \pi N$
1595	ARNDT	06	DPWA $\pi N \rightarrow \pi N$ , $\eta N$
1608	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1600 $\pm$ 15	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1597 $\pm$ 4	ANISOVICH	12A	DPWA Multichannel
1587	SHRESTHA	12A	DPWA Multichannel
1607	VRANA	00	DPWA Multichannel

–2 <i>x</i> IMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>120 to 140 (<math>\approx 130</math>) OUR ESTIMATE</b>			
134 $\pm$ 8	SOKHOYAN	15A	DPWA Multichannel
114 $\pm$ 12 $\pm$ 4	SVARC	14	L+P $\pi N \rightarrow \pi N$
135	ARNDT	06	DPWA $\pi N \rightarrow \pi N$ , $\eta N$
116	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
120 $\pm$ 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
130 $\pm$ 9	ANISOVICH	12A	DPWA Multichannel
107	SHRESTHA	12A	DPWA Multichannel
148	VRANA	00	DPWA Multichannel

$\Delta(1620)$ ELASTIC POLE RESIDUE			
MODULUS $ r $			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>15 to 20 (<math>\approx 17</math>) OUR ESTIMATE</b>			
20 $\pm$ 3	SOKHOYAN	15A	DPWA Multichannel
17 $\pm$ 2 $\pm$ 1	SVARC	14	L+P $\pi N \rightarrow \pi N$
15	ARNDT	06	DPWA $\pi N \rightarrow \pi N$ , $\eta N$
19	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
15 $\pm$ 2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
18 $\pm$ 2	ANISOVICH	12A	DPWA Multichannel

PHASE $\theta$			
VALUE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
<b>– 90 to –110 (<math>\approx -100</math>) OUR ESTIMATE</b>			
– 90 $\pm$ 15	SOKHOYAN	15A	DPWA Multichannel
–106 $\pm$ 10 $\pm$ 4	SVARC	14	L+P $\pi N \rightarrow \pi N$
– 92	ARNDT	06	DPWA $\pi N \rightarrow \pi N$ , $\eta N$
– 95	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
–110 $\pm$ 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–100 $\pm$ 5	ANISOVICH	12A	DPWA Multichannel

# Baryon Particle Listings

## $\Delta(1620)$ , $\Delta(1700)$

### $\Delta(1620)$ INELASTIC POLE RESIDUE

The "normalized residue" is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow \Delta(1620) \rightarrow \Delta\pi$ , $D$ -wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.42±0.06	−90 ± 20	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.38±0.09	−85 ± 30	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1620) \rightarrow N(1440)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.10±0.06	−65 ± 30	SOKHOYAN	15A	DPWA Multichannel

### $\Delta(1620)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1600 to 1660 (≈ 1630) OUR ESTIMATE</b>			
1595 ± 8	SOKHOYAN	15A	DPWA Multichannel
1615.2± 0.4	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1620 ± 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1610 ± 7	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1600 ± 8	ANISOVICH	12A	DPWA Multichannel
1600 ± 1	SHRESTHA	12A	DPWA Multichannel
1612 ± 2	PENNER	02c	DPWA Multichannel
1617 ± 15	VRANA	00	DPWA Multichannel

### $\Delta(1620)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>130 to 150 (≈ 140) OUR ESTIMATE</b>			
135 ± 9	SOKHOYAN	15A	DPWA Multichannel
146.9± 1.9	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
140 ± 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
139 ± 18	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
130 ± 11	ANISOVICH	12A	DPWA Multichannel
112 ± 2	SHRESTHA	12A	DPWA Multichannel
202 ± 7	PENNER	02c	DPWA Multichannel
143 ± 42	VRANA	00	DPWA Multichannel

### $\Delta(1620)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	20–30 %
$\Gamma_2$ $N\pi\pi$	55–80 %
$\Gamma_3$ $\Delta(1232)\pi$	
$\Gamma_4$ $\Delta(1232)\pi$ , $D$ -wave	52–72 %
$\Gamma_5$ $N\rho$	
$\Gamma_6$ $N\rho$ , $S=1/2$ , $S$ -wave	seen
$\Gamma_7$ $N\rho$ , $S=3/2$ , $D$ -wave	seen
$\Gamma_8$ $N(1440)\pi$	3–9 %
$\Gamma_9$ $N\gamma$ , helicity=1/2	0.03–0.10 %

### $\Delta(1620)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>20 to 30 OUR ESTIMATE</b>				
28 ± 3	SOKHOYAN	15A	DPWA Multichannel	
31.5±0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$	
25 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
35 ± 6	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
28 ± 3	ANISOVICH	12A	DPWA Multichannel	
33 ± 2	SHRESTHA	12A	DPWA Multichannel	
34 ± 1	PENNER	02c	DPWA Multichannel	
45 ± 5	VRANA	00	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
<b>62±10</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
60±17	ANISOVICH	12A	DPWA Multichannel	
32± 2	SHRESTHA	12A	DPWA Multichannel	
39± 2	VRANA	00	DPWA Multichannel	

$\Gamma(N\rho, S=1/2, S\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
<b>26±2</b>				
14±3	SHRESTHA	12A	DPWA Multichannel	
	VRANA	00	DPWA Multichannel	

$\Gamma(N\rho, S=3/2, D\text{-wave})/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
<b>2±1</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
	VRANA	00	DPWA Multichannel	

$\Gamma(N(1440)\pi)/\Gamma_{total}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
<b>6±3</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
9±1	SHRESTHA	12A	DPWA Multichannel	
0±1	VRANA	00	DPWA Multichannel	

### $\Delta(1620)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $\Delta(1620) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.054±0.007	−6 ± 7	SOKHOYAN	15A	DPWA Multichannel

### $\Delta(1620)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $\Delta(1620) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>+0.040±0.015 OUR ESTIMATE</b>			
0.055±0.007	SOKHOYAN	15A	DPWA Multichannel
0.029±0.003	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$
0.050±0.002	DUGGER	07	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.052±0.005	ANISOVICH	12A	DPWA Multichannel
−0.003±0.003	SHRESTHA	12A	DPWA Multichannel
0.066	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$
−0.050	PENNER	02D	DPWA Multichannel

### $\Delta(1620)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $\Delta(1620)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER	07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	79	$\pi N$ Newsletter 9 1	G. H��hler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
HOEHLER	79	PDAT 12-1	G. H��hler <i>et al.</i>	(KARLT) IUP
Also		Toronto Conf. 3	R. Koch	(KARLT) IUP

## $\Delta(1700) 3/2^-$

$I(J^P) = \frac{3}{2}(\frac{3}{2}^-)$  Status: \*\*\*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $\Delta(1700)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1620 to 1680 (≈ 1650) OUR ESTIMATE</b>			
1685±10	SOKHOYAN	15A	DPWA Multichannel
1643± 6±3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1632	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1651	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1675±25	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1685±10	GUTZ	14	DPWA Multichannel
1680±10	ANISOVICH	12A	DPWA Multichannel
1656	SHRESTHA	12A	DPWA Multichannel
1726	VRANA	00	DPWA Multichannel

See key on page 601

Baryon Particle Listings  
 $\Delta(1700)$

-2×IMAGINARY PART				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
160 to 300 (≈ 230) OUR ESTIMATE				
300±15	SOKHOYAN	15A	DPWA	Multichannel
217±10±8	<sup>1</sup> SVARC	14	L+P	$\pi N \rightarrow \pi N$
253	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
159	HOEHLER	93	SPED	$\pi N \rightarrow \pi N$
220±40	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
300±15	GUTZ	14	DPWA	Multichannel
305±15	ANISOVICH	12A	DPWA	Multichannel
226	SHRESTHA	12A	DPWA	Multichannel
118	VRANA	00	DPWA	Multichannel

$\Delta(1700)$ ELASTIC POLE RESIDUE				
MODULUS $ r $				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
10 to 40 ( $\approx$ 25) OUR ESTIMATE				
40 $\pm$ 6	1	SOKHOYAN	15A	DPWA Multichannel
13 $\pm$ 1 $\pm$ 1		SVARC	14	L+P $\pi N \rightarrow \pi N$
18		ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
10		HOEHLER	93	SPED $\pi N \rightarrow \pi N$
13 $\pm$ 3		CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
40 $\pm$ 6		GUTZ	14	DPWA Multichannel
42 $\pm$ 7		ANISOVICH	12A	DPWA Multichannel

PHASE $\theta$				
VALUE (°)	DOCUMENT ID	TECN	COMMENT	
<b>−40 to 0 (≈ − 20) OUR ESTIMATE</b>				
− 1±10	SOKHOYAN	15A	DPWA	Multichannel
−30± 4±3	<sup>1</sup> SVARC	14	L+P	$\pi N \rightarrow \pi N$
−40	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
−20±25	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
− 1±10	GUTZ	14	DPWA	Multichannel
− 3±15	ANISOVICH	12A	DPWA	Multichannel

$\Delta(1700)$ INELASTIC POLE RESIDUE				
The “normalized residue” is the residue divided by $\Gamma_{pole}/2$ .				
Normalized residue in $N\pi \rightarrow \Delta(1700) \rightarrow \Delta\eta$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.12±0.02	−60±12	GUTZ	14	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.12±0.03	−60±15	ANISOVICH	12A	DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1700) \rightarrow N(1535)\pi$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.035±0.015	−75±30	GUTZ	14	DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1700) \rightarrow \Delta(1232)\pi, S\text{-wave}$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.25±0.12	135±45	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1700) \rightarrow \Delta(1232)\pi, D\text{-wave}$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.12±0.06	−160±30	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1700) \rightarrow N(1520)\pi, P\text{-wave}$				
MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.10±0.03	−10±20	SOKHOYAN	15A	DPWA Multichannel

$\Delta(1700)$ BREIT-WIGNER MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
1670 to 1750 ( $\approx 1700$ ) OUR ESTIMATE				
1715 $\pm 20$	SOKHOYAN	15A	DPWA	Multichannel
1695.0 $\pm$ 1.3	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
1710 $\pm 30$	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
1680 $\pm 70$	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1715 $\pm 20$	GUTZ	14	DPWA	Multichannel
1715 $\begin{smallmatrix} +30 \\ -15 \end{smallmatrix}$	ANISOVICH	12A	DPWA	Multichannel
1691 $\pm 4$	SHRESTHA	12A	DPWA	Multichannel
1678 $\pm 1$	PENNER	02c	DPWA	Multichannel
1732 $\pm 23$	VRA NA	00	DPWA	Multichannel

$\Delta(1700)$ BREIT-WIGNER WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
200 to 400 ( $\approx$ 300) OUR ESTIMATE				
300 $\pm$ 25	SOKHOYAN	15A	DPWA	Multichannel
375.5 $\pm$ 7.0	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
280 $\pm$ 80	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
230 $\pm$ 80	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •				
300 ±25	GUTZ	14	DPWA	Multichannel
310 +40 −15	ANISOVICH	12A	DPWA	Multichannel
248 ±9	SHRESTHA	12A	DPWA	Multichannel
606 ±15	PENNER	02c	DPWA	Multichannel
119 ±70	VRANA	00	DPWA	Multichannel

$\Delta(1700)$ DECAY MODES		
The following branching fractions are our estimates, not fits or averages.		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\pi$	10–20 %	
$\Gamma_2$ $N\pi\pi$	10–55 %	
$\Gamma_3$ $\Delta(1232)\pi$	10–50 %	
$\Gamma_4$ $\Delta(1232)\pi, S\text{-wave}$	5–35 %	
$\Gamma_5$ $\Delta(1232)\pi, D\text{-wave}$	4–16 %	
$\Gamma_6$ $N\rho$		
$\Gamma_7$ $N\rho, S=3/2, S\text{-wave}$	seen	
$\Gamma_8$ $N(1520)\pi, P\text{-wave}$	1–5 %	
$\Gamma_9$ $N(1535)\pi$	0.5–1.5 %	
$\Gamma_{10}$ $\Delta(1232)\eta$	3–7 %	
$\Gamma_{11}$ $N\gamma$	0.22–0.60 %	
$\Gamma_{12}$ $N\gamma, \text{helicity}=1/2$	0.12–0.30 %	
$\Gamma_{13}$ $N\gamma, \text{helicity}=3/2$	0.10–0.30 %	

$\Delta(1700)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)	DOCUMENT ID	TECN	COMMENT
10 to 20 OUR ESTIMATE			
22 ±4	SOKHOYAN	15A	DPWA Multichannel
15.6±0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
12 ±3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
20 ±3	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
22 ±4	GUTZ	14	DPWA Multichannel
22 ±4	ANISOVICH	12A	DPWA Multichannel
14 ±1	SHRESTHA	12A	DPWA Multichannel
14 ±1	PENNER	02c	DPWA Multichannel
5 ±1	VRANA	00	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, S\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
20±15	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
20+ <sup>25</sup> −13	ANISOVICH	12A	DPWA Multichannel
54±3	SHRESTHA	12A	DPWA Multichannel
90±2	VRANA	00	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
10±6	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
12+ <sup>14</sup> −7	ANISOVICH	12A	DPWA Multichannel
1±1	SHRESTHA	12A	DPWA Multichannel
4±1	VRANA	00	DPWA Multichannel

$\Gamma(N\rho, S=3/2, S\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
30±3	SHRESTHA	12A	DPWA Multichannel
1±1	VRANA	00	DPWA Multichannel

$\Gamma(N(1520)\pi, P\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
3±2	SOKHOYAN	15A	DPWA Multichannel

$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
1.0±0.5	GUTZ	14	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4 ±2	HORN	08A	DPWA Multichannel

$\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
5±2	GUTZ	14	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5±2	ANISOVICH	12A	DPWA Multichannel

Baryon Particle Listings

$\Delta(1700)$ ,  $\Delta(1750)$

$\Gamma(N(1535)\pi)/\Gamma(\Delta(1232)\eta)$				$\Gamma_9/\Gamma_{10}$
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.67	KASHEVAROV 09	CBAL	$\gamma p \rightarrow p \pi^0 \eta$	

$\Delta(1700)$  PHOTON DECAY AMPLITUDES AT THE POLE

$\Delta(1700) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$					
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT	
0.175 $\pm$ 0.020	50 $\pm$ 10	SOKHOYAN	15A	DPWA	Multichannel

$\Delta(1700) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$					
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT	
0.180 $\pm$ 0.020	45 $\pm$ 10	SOKHOYAN	15A	DPWA	Multichannel

$\Delta(1700)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES

$\Delta(1700) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$					
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT		
<b>0.140<math>\pm</math>0.030 OUR ESTIMATE</b>					
0.165 $\pm$ 0.020	SOKHOYAN	15A	DPWA	Multichannel	
0.132 $\pm$ 0.005	DUGGER	13	DPWA	$\gamma N \rightarrow \pi N$	
0.105 $\pm$ 0.005	WORKMAN	12A	DPWA	$\gamma N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.165 $\pm$ 0.020	GUTZ	14	DPWA	Multichannel	
0.160 $\pm$ 0.020	ANISOVICH	12A	DPWA	Multichannel	
0.058 $\pm$ 0.010	SHRESTHA	12A	DPWA	Multichannel	
0.226	DRECHSEL	07	DPWA	$\gamma N \rightarrow \pi N$	
0.125 $\pm$ 0.003	DUGGER	07	DPWA	$\gamma N \rightarrow \pi N$	
0.096	PENNER	02D	DPWA	Multichannel	

$\Delta(1700) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$					
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT		
<b>0.140<math>\pm</math>0.030 OUR ESTIMATE</b>					
0.170 $\pm$ 0.025	SOKHOYAN	15A	DPWA	Multichannel	
0.108 $\pm$ 0.005	DUGGER	13	DPWA	$\gamma N \rightarrow \pi N$	
0.092 $\pm$ 0.004	WORKMAN	12A	DPWA	$\gamma N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.170 $\pm$ 0.025	GUTZ	14	DPWA	Multichannel	
0.165 $\pm$ 0.025	ANISOVICH	12A	DPWA	Multichannel	
0.097 $\pm$ 0.008	SHRESTHA	12A	DPWA	Multichannel	
0.210	DRECHSEL	07	DPWA	$\gamma N \rightarrow \pi N$	
0.105 $\pm$ 0.003	DUGGER	07	DPWA	$\gamma N \rightarrow \pi N$	
0.154	PENNER	02D	DPWA	Multichannel	

$\Delta(1700)$  FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

$\Delta(1700)$  REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN 15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ 14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG 14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC 14	PR C89 045205	A. Svarc <i>et al.</i>	
DUGGER 13	PR C88 065203	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ANISOVICH 12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA 12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN 12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
KASHEVAROV 09	EPJ A42 141	V.L. Kashevarov <i>et al.</i>	(MAMI Crystal Ball/TAPS)
HORN 08A	EPJ A38 173	I. Horn <i>et al.</i>	(CB-ELSA Collab.)
Also	PRL 101 202002	I. Horn <i>et al.</i>	(CB-ELSA Collab.)
DRECHSEL 07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
DUGGER 07	PR C76 025211	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ARNDT 06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
PENNER 02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER 02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA 00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER 93	$\pi N$ Newsletter 9 1	G. H�hler	(KARL)
CUTKOSKY 80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also	PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER 79	PDAT 12-1	G. H�hler <i>et al.</i>	(KARLT) IJP
Also	Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(1750) 1/2^+$

$I(J^P) = \frac{3}{2}(\frac{1}{2}^+)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Delta(1750)$  POLE POSITION

REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1748	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1714	VRANA	00	DPWA Multichannel

$-\text{2}\times\text{IMAGINARY PART}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
524	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
68	VRANA	00	DPWA Multichannel

$\Delta(1750)$  ELASTIC POLE RESIDUE

MODULUS $ r $			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
48	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$

PHASE $\theta$			
VALUE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
158	ARNDT	04	DPWA $\pi N \rightarrow \pi N, \eta N$

$\Delta(1750)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1712 $\pm$ 1	PENNER	02c	DPWA Multichannel
1721 $\pm$ 61	VRANA	00	DPWA Multichannel

$\Delta(1750)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
643 $\pm$ 17	PENNER	02c	DPWA Multichannel
70 $\pm$ 50	VRANA	00	DPWA Multichannel

$\Delta(1750)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	seen
$\Gamma_2$ $N\pi\pi$	
$\Gamma_3$ $N(1440)\pi$	seen
$\Gamma_4$ $\Sigma K$	seen

$\Delta(1750)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1 $\pm$ 1	PENNER	02c	DPWA Multichannel	
6 $\pm$ 9	VRANA	00	DPWA Multichannel	

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
83 $\pm$ 1	VRANA	00	DPWA Multichannel	

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.1 $\pm$ 0.1	PENNER	02c	DPWA Multichannel	

$\Delta(1750)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES

Papers on  $\gamma N$  amplitudes predating 1981 may be found in our 2006 edition, Journal of Physics **G33** 1 (2006).

$\Delta(1750) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$					
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.053	PENNER	02D	DPWA	Multichannel	

$\Delta(1750)$  REFERENCES

PDG 06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
ARNDT 04	PR C69 035213	R.A. Arndt <i>et al.</i>	(GWU, TRIU)
PENNER 02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER 02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA 00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)



See key on page 601

Baryon Particle Listings  
 $\Delta(1900)$

$\Delta(1900) \ 1/2^-$

$I(J^P) = \frac{3}{2}(\frac{1}{2}^-)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

$\Delta(1900)$ POLE POSITION			
REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1845 ± 20	SOKHOYAN	15A	DPWA Multichannel
1865 ± 35 ± 19	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1780	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1870 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1845 ± 20	GUTZ	14	DPWA Multichannel
1845 ± 25	ANISOVICH	12A	DPWA Multichannel
1844	SHRESTHA	12A	DPWA Multichannel
1795	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
295 ± 35	SOKHOYAN	15A	DPWA Multichannel
187 ± 50 ± 19	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
180 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
295 ± 35	GUTZ	14	DPWA Multichannel
300 ± 45	ANISOVICH	12A	DPWA Multichannel
223	SHRESTHA	12A	DPWA Multichannel
58	VRANA	00	DPWA Multichannel

$\Delta(1900)$ ELASTIC POLE RESIDUE			
MODULUS $ r $			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
11 ± 2	SOKHOYAN	15A	DPWA Multichannel
11 ± 4 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
10 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
11 ± 2	GUTZ	14	DPWA Multichannel
10 ± 3	ANISOVICH	12A	DPWA Multichannel

PHASE $\theta$			
VALUE (°)	DOCUMENT ID	TECN	COMMENT
−115 ± 20	SOKHOYAN	15A	DPWA Multichannel
20 ± 27 ± 19	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
+ 20 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−115 ± 20	GUTZ	14	DPWA Multichannel
−125 ± 20	ANISOVICH	12A	DPWA Multichannel

$\Delta(1900)$ INELASTIC POLE RESIDUE			
The “normalized residue” is the residue divided by $\Gamma_{pole}/2$ .			

Normalized residue in $N\pi \rightarrow \Delta(1900) \rightarrow \Sigma K$			
MODULUS	PHASE (°)	DOCUMENT ID	TECN COMMENT
0.07 ± 0.02	−50 ± 30	ANISOVICH	12A DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1900) \rightarrow \Delta\pi, D\text{-wave}$			
MODULUS	PHASE (°)	DOCUMENT ID	TECN COMMENT
0.18 ± 0.10	105 ± 25	SOKHOYAN	15A DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.12 <sup>+0.08</sup> <sub>−0.05</sub>	110 ± 20	ANISOVICH	12A DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1900) \rightarrow \Delta(1232)\eta$			
MODULUS	PHASE (°)	DOCUMENT ID	TECN COMMENT
0.013 ± 0.006	undefined	GUTZ	14 DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1900) \rightarrow N(1440)\pi$			
MODULUS	PHASE (°)	DOCUMENT ID	TECN COMMENT
0.11 ± 0.06	115 ± 30	SOKHOYAN	15A DPWA Multichannel

Normalized residue in $N\pi \rightarrow \Delta(1900) \rightarrow N(1520)\pi$			
MODULUS	PHASE (°)	DOCUMENT ID	TECN COMMENT
0.06 ± 0.03	undefined	SOKHOYAN	15A DPWA Multichannel

$\Delta(1900)$ BREIT-WIGNER MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1840 to 1920 (≈ 1860) OUR ESTIMATE			
1840 ± 20	SOKHOYAN	15A	DPWA Multichannel
1890 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1908 ± 30	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •			
1840 ± 20	GUTZ	14	DPWA Multichannel
1840 ± 30	ANISOVICH	12A	DPWA Multichannel
1868 ± 12	SHRESTHA	12A	DPWA Multichannel
1802 ± 87	VRANA	00	DPWA Multichannel

$\Delta(1900)$ BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
295 ± 30	SOKHOYAN	15A	DPWA Multichannel
170 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
140 ± 40	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
295 ± 30	GUTZ	14	DPWA Multichannel
300 ± 45	ANISOVICH	12A	DPWA Multichannel
234 ± 27	SHRESTHA	12A	DPWA Multichannel
48 ± 45	VRANA	00	DPWA Multichannel

$\Delta(1900)$ DECAY MODES		
The following branching fractions are our estimates, not fits or averages.		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1 \ N\pi$	4–12 %	
$\Gamma_2 \ \Sigma K$	seen	
$\Gamma_3 \ N\pi\pi$	45–85 %	
$\Gamma_4 \ \Delta(1232)\pi$		
$\Gamma_5 \ \Delta(1232)\pi, D\text{-wave}$	30–70 %	
$\Gamma_6 \ N\rho$		
$\Gamma_7 \ N\rho, S=1/2, S\text{-wave}$	seen	
$\Gamma_8 \ N\rho, S=3/2, D\text{-wave}$	seen	
$\Gamma_9 \ N(1440)\pi$	8–32 %	
$\Gamma_{10} \ N(1520)\pi$	2–10 %	
$\Gamma_{11} \ \Delta(1232)\eta$	0–2 %	
$\Gamma_{12} \ N\gamma, \text{ helicity}=1/2$	0.06–0.43 %	

$\Delta(1900)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT
7 ± 2	SOKHOYAN	15A	DPWA Multichannel
10 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
8 ± 4	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
7 ± 2	GUTZ	14	DPWA Multichannel
7 ± 3	ANISOVICH	12A	DPWA Multichannel
8 ± 1	SHRESTHA	12A	DPWA Multichannel
33 ± 10	VRANA	00	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
50 ± 20	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
15 <sup>+50</sup> <sub>−10</sub>	ANISOVICH	12A	DPWA Multichannel
56 ± 6	SHRESTHA	12A	DPWA Multichannel
28 ± 1	VRANA	00	DPWA Multichannel

$\Gamma(N\rho, S=1/2, S\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
12 ± 4	SHRESTHA	12A	DPWA Multichannel
30 ± 2	VRANA	00	DPWA Multichannel

$\Gamma(N\rho, S=3/2, D\text{-wave})/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
23 ± 5	SHRESTHA	12A	DPWA Multichannel
5 ± 1	VRANA	00	DPWA Multichannel

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
20 ± 12	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
< 1	SHRESTHA	12A	DPWA Multichannel
4 ± 1	VRANA	00	DPWA Multichannel

$\Gamma(N(1520)\pi)/\Gamma_{\text{total}}$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT
6 ± 4	SOKHOYAN	15A	DPWA Multichannel

# Baryon Particle Listings

## $\Delta(1900)$ , $\Delta(1905)$

$\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}$				$\Gamma_{11}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
1±1	GUTZ	14	DPWA	Multichannel

### $\Delta(1900)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $\Delta(1900) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.064±0.015	60 ± 20	SOKHOYAN	15A	DPWA Multichannel

### $\Delta(1900)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $\Delta(1900) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.065 ± 0.015	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.057 ± 0.014	GUTZ	14	DPWA Multichannel
−0.082 ± 0.009	SHRESTHA	12A	DPWA Multichannel

### $\Delta(1900)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $\Delta(1900)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

## $\Delta(1905)$ 5/2<sup>+</sup>

$$I(J^P) = \frac{3}{2}(\frac{5}{2}^+) \text{ Status: } ****$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $\Delta(1905)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1805 to 1835 (≈ 1820) OUR ESTIMATE</b>			
1800 ± 6	SOKHOYAN	15A	DPWA Multichannel
1752 ± 3 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1819	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1829	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1830 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1800 ± 6	GUTZ	14	DPWA Multichannel
1805 ± 10	ANISOVICH	12A	DPWA Multichannel
1769	SHRESTHA	12A	DPWA Multichannel
1793	VRANA	00	DPWA Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>265 to 300 (≈ 280) OUR ESTIMATE</b>			
290 ± 15	SOKHOYAN	15A	DPWA Multichannel
346 ± 6 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
247	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
303	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
280 ± 60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
290 ± 15	GUTZ	14	DPWA Multichannel
300 ± 15	ANISOVICH	12A	DPWA Multichannel
239	SHRESTHA	12A	DPWA Multichannel
302	VRANA	00	DPWA Multichannel

### $\Delta(1905)$ ELASTIC POLE RESIDUE

#### MODULUS |r|

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>15 to 25 (≈ 20) OUR ESTIMATE</b>			
19 ± 2	SOKHOYAN	15A	DPWA Multichannel
24 ± 1 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
15	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
25	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
25 ± 8	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
19 ± 2	GUTZ	14	DPWA Multichannel
20 ± 2	ANISOVICH	12A	DPWA Multichannel

### PHASE $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
<b>−120 to −30 (≈ −50) OUR ESTIMATE</b>			
−45 ± 4	SOKHOYAN	15A	DPWA Multichannel
−114 ± 1 ± 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−30	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−50 ± 20	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−45 ± 4	GUTZ	14	DPWA Multichannel
−44 ± 5	ANISOVICH	12A	DPWA Multichannel

### $\Delta(1905)$ INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{\text{pole}}/2$ .

#### Normalized residue in $N\pi \rightarrow \Delta(1905) \rightarrow \Delta\pi$ , $P$ -wave

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
19 ± 7	10 ± 30	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
25 ± 6	0 ± 15	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1905) \rightarrow N(1535)\pi$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
2.5 ± 1.0	130 ± 35	GUTZ	14	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1905) \rightarrow \Delta(1232)\eta$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
7 ± 2	40 ± 20	GUTZ	14	DPWA Multichannel

### $\Delta(1905)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1855 to 1910 (≈ 1880) OUR ESTIMATE</b>			
1856 ± 6	SOKHOYAN	15A	DPWA Multichannel
1857.8 ± 1.6	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1910 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1905 ± 20	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1856 ± 6	GUTZ	14	DPWA Multichannel
1861 ± 6	ANISOVICH	12A	DPWA Multichannel
1818 ± 8	SHRESTHA	12A	DPWA Multichannel
1873 ± 77	VRANA	00	DPWA Multichannel

### $\Delta(1905)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>270 to 400 (≈ 330) OUR ESTIMATE</b>			
325 ± 15	SOKHOYAN	15A	DPWA Multichannel
320.6 ± 8.6	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
400 ± 100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
260 ± 20	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
325 ± 15	GUTZ	14	DPWA Multichannel
335 ± 18	ANISOVICH	12A	DPWA Multichannel
278 ± 18	SHRESTHA	12A	DPWA Multichannel
461 ± 111	VRANA	00	DPWA Multichannel

### $\Delta(1905)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	9–15 %
$\Gamma_2$ $N\pi\pi$	
$\Gamma_3$ $\Delta(1232)\pi$	
$\Gamma_4$ $\Delta(1232)\pi$ , $P$ -wave	23–43 %
$\Gamma_5$ $\Delta(1232)\pi$ , $F$ -wave	seen
$\Gamma_6$ $N\rho$	
$\Gamma_7$ $N\rho$ , $S=3/2$ , $P$ -wave	seen
$\Gamma_8$ $N(1535)\pi$	< 1 %
$\Gamma_9$ $N(1680)\pi$ , $P$ -wave	5–15 %
$\Gamma_{10}$ $\Delta(1232)\eta$	2–6 %
$\Gamma_{11}$ $N\gamma$	0.012–0.036 %
$\Gamma_{12}$ $N\gamma$ , helicity=1/2	0.002–0.006 %
$\Gamma_{13}$ $N\gamma$ , helicity=3/2	0.01–0.03 %

See key on page 601

# Baryon Particle Listings

## $\Delta(1905)$ , $\Delta(1910)$

 **$\Delta(1905)$  BRANCHING RATIOS**

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
<b>9 to 15 OUR ESTIMATE</b>				
13 $\pm$ 2	SOKHOYAN	15A	DPWA Multichannel	
12.2 $\pm$ 0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$	
8 $\pm$ 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
15 $\pm$ 2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13 $\pm$ 2	GUTZ	14	DPWA Multichannel	
13 $\pm$ 2	ANISOVICH	12A	DPWA Multichannel	
6 $\pm$ 1	SHRESTHA	12A	DPWA Multichannel	
9 $\pm$ 1	VRANA	00	DPWA Multichannel	
<b><math>\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}</math></b>				$\Gamma_4/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
33 $\pm$ 10	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
45 $\pm$ 14	ANISOVICH	12A	DPWA Multichannel	
28 $\pm$ 7	SHRESTHA	12A	DPWA Multichannel	
23 $\pm$ 1	VRANA	00	DPWA Multichannel	
<b><math>\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}</math></b>				$\Gamma_5/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
64 $\pm$ 8	SHRESTHA	12A	DPWA Multichannel	
44 $\pm$ 1	VRANA	00	DPWA Multichannel	
<b><math>\Gamma(N\rho, S=3/2, P\text{-wave})/\Gamma_{\text{total}}</math></b>				$\Gamma_7/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 6	SHRESTHA	12A	DPWA Multichannel	
24 $\pm$ 1	VRANA	00	DPWA Multichannel	
<b><math>\Gamma(N(1535)\pi)/\Gamma_{\text{total}}</math></b>				$\Gamma_8/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<1	GUTZ	14	DPWA Multichannel	
<b><math>\Gamma(N(1680)\pi, P\text{-wave})/\Gamma_{\text{total}}</math></b>				$\Gamma_9/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
10 $\pm$ 5	SOKHOYAN	15A	DPWA Multichannel	
<b><math>\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}</math></b>				$\Gamma_{10}/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
4 $\pm$ 2	GUTZ	14	DPWA Multichannel	

 **$\Delta(1905)$  PHOTON DECAY AMPLITUDES AT THE POLE** **$\Delta(1905) \rightarrow N\gamma$ , helicity-1/2 amplitude  $A_{1/2}$** 

MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
0.025 $\pm$ 0.005	-28 $\pm$ 12	SOKHOYAN	15A	DPWA Multichannel

 **$\Delta(1905) \rightarrow N\gamma$ , helicity-3/2 amplitude  $A_{3/2}$** 

MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
-0.050 $\pm$ 0.004	5 $\pm$ 10	SOKHOYAN	15A	DPWA Multichannel

 **$\Delta(1905)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES** **$\Delta(1905) \rightarrow N\gamma$ , helicity-1/2 amplitude  $A_{1/2}$** 

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>+0.022<math>\pm</math>0.005 OUR ESTIMATE</b>			
0.025 $\pm$ 0.005	SOKHOYAN	15A	DPWA Multichannel
0.020 $\pm$ 0.002	DUGGER	13	DPWA $\gamma N \rightarrow \pi N$
0.019 $\pm$ 0.002	WORKMAN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.025 $\pm$ 0.005	GUTZ	14	DPWA Multichannel
0.025 $\pm$ 0.004	ANISOVICH	12A	DPWA Multichannel
0.066 $\pm$ 0.018	SHRESTHA	12A	DPWA Multichannel
0.018	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

 **$\Delta(1905) \rightarrow N\gamma$ , helicity-3/2 amplitude  $A_{3/2}$** 

VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT
<b>-0.045<math>\pm</math>0.010 OUR ESTIMATE</b>			
-0.050 $\pm$ 0.005	SOKHOYAN	15A	DPWA Multichannel
-0.049 $\pm$ 0.005	DUGGER	13	DPWA $\gamma N \rightarrow \pi N$
-0.038 $\pm$ 0.004	WORKMAN	12A	DPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.050 $\pm$ 0.005	GUTZ	14	DPWA Multichannel
-0.049 $\pm$ 0.004	ANISOVICH	12A	DPWA Multichannel
-0.223 $\pm$ 0.029	SHRESTHA	12A	DPWA Multichannel
-0.028	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$

 **$\Delta(1905)$  FOOTNOTES**<sup>1</sup> Fit to the amplitudes of HOEHLER 79. **$\Delta(1905)$  REFERENCES**For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
DUGGER	13	PR C88 065203	M. Dugger <i>et al.</i>	(JLab CLAS Collab.)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman <i>et al.</i>	(GWU)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	79	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

 **$\Delta(1910) 1/2^+$** 

$$I(J^P) = \frac{3}{2}(\frac{1}{2}^+) \text{ Status: } ***$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014). **$\Delta(1910)$  POLE POSITION****REAL PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1830 to 1880 (<math>\approx</math> 1855) OUR ESTIMATE</b>			
1840 $\pm$ 40	SOKHOYAN	15A	DPWA Multichannel
1896 $\pm$ 11	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1771	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1874	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1880 $\pm$ 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1840 $\pm$ 40	GUTZ	14	DPWA Multichannel
1850 $\pm$ 40	ANISOVICH	12A	DPWA Multichannel
1910	SHRESTHA	12A	DPWA Multichannel
1880	VRANA	00	DPWA Multichannel

**-2 $\times$ IMAGINARY PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>200 to 500 (<math>\approx</math> 350) OUR ESTIMATE</b>			
370 $\pm$ 60	SOKHOYAN	15A	DPWA Multichannel
302 $\pm$ 22	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
479	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
283	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
200 $\pm$ 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
370 $\pm$ 60	GUTZ	14	DPWA Multichannel
350 $\pm$ 45	ANISOVICH	12A	DPWA Multichannel
199	SHRESTHA	12A	DPWA Multichannel
496	VRANA	00	DPWA Multichannel

 **$\Delta(1910)$  ELASTIC POLE RESIDUE****MODULUS  $|r|$** 

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>20 to 45 (<math>\approx</math> 30) OUR ESTIMATE</b>			
25 $\pm$ 6	SOKHOYAN	15A	DPWA Multichannel
29 $\pm$ 2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
45	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
38	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
20 $\pm$ 4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
25 $\pm$ 6	GUTZ	14	DPWA Multichannel
24 $\pm$ 6	ANISOVICH	12A	DPWA Multichannel

**PHASE  $\theta$** 

VALUE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
<b>-80 to -180 (<math>\approx</math> -130) OUR ESTIMATE</b>			
-155 $\pm$ 30	SOKHOYAN	15A	DPWA Multichannel
-83 $\pm$ 4 $\pm$ 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
+172	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
-90 $\pm$ 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-155 $\pm$ 30	GUTZ	14	DPWA Multichannel
-145 $\pm$ 30	ANISOVICH	12A	DPWA Multichannel

# Baryon Particle Listings

## $\Delta(1910)$ , $\Delta(1920)$

### $\Delta(1910)$ INELASTIC POLE RESIDUE

The "normalized residue" is the residue divided by  $\Gamma_{pole}/2$ .

#### Normalized residue in $N\pi \rightarrow \Delta(1910) \rightarrow \Sigma K$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.07±0.02	−110 ± 30	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1910) \rightarrow \Delta\pi$ , $P$ -wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.24±0.10	85 ± 35	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.16±0.09	95 ± 40	ANISOVICH	12A	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1910) \rightarrow \Delta(1232)\eta$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.11±0.04	−150 ± 50	GUTZ	14	DPWA Multichannel

#### Normalized residue in $N\pi \rightarrow \Delta(1910) \rightarrow N(1440)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.06±0.03	170 ± 45	SOKHOYAN	15A	DPWA Multichannel

### $\Delta(1910)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1860 to 1910 (≈ 1890) OUR ESTIMATE</b>			
1845 ± 40	SOKHOYAN	15A	DPWA Multichannel
2067.9 ± 1.7	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1910 ± 40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1888 ± 20	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1845 ± 40	GUTZ	14	DPWA Multichannel
1860 ± 40	ANISOVICH	12A	DPWA Multichannel
1934 ± 5	SHRESTHA	12A	DPWA Multichannel
1995 ± 12	VRANA	00	DPWA Multichannel

### $\Delta(1910)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>220 to 340 (≈ 280) OUR ESTIMATE</b>			
360 ± 60	SOKHOYAN	15A	DPWA Multichannel
543 ± 10	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
225 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
280 ± 50	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
360 ± 60	GUTZ	14	DPWA Multichannel
350 ± 55	ANISOVICH	12A	DPWA Multichannel
211 ± 11	SHRESTHA	12A	DPWA Multichannel
713±465	VRANA	00	DPWA Multichannel

### $\Delta(1910)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	15–30 %
$\Gamma_2$ $\Sigma K$	4–14 %
$\Gamma_3$ $N\pi\pi$	
$\Gamma_4$ $\Delta(1232)\pi$	34–66 %
$\Gamma_5$ $N(1440)\pi$	3–9 %
$\Gamma_6$ $\Delta(1232)\eta$	5–13 %
$\Gamma_7$ $N\gamma$ , helicity=1/2	0.0–0.02 %

### $\Delta(1910)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>15 to 30 OUR ESTIMATE</b>				
12 ± 3	SOKHOYAN	15A	DPWA Multichannel	
23.9± 0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$	
19 ± 3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
24 ± 6	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
12 ± 3	GUTZ	14	DPWA Multichannel	
12 ± 3	ANISOVICH	12A	DPWA Multichannel	
17 ± 1	SHRESTHA	12A	DPWA Multichannel	
29 ± 21	VRANA	00	DPWA Multichannel	

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
<b>VALUE (%)</b>				
9±5	ANISOVICH	12A	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
<b>VALUE (%)</b>				
50±16	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
60±28	ANISOVICH	12A	DPWA Multichannel	

$\Gamma(N(1440)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
<b>VALUE (%)</b>				
6±3	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
47±6	SHRESTHA	12A	DPWA Multichannel	
56±7	VRANA	00	DPWA Multichannel	

$\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
<b>VALUE (%)</b>				
9±4	GUTZ	14	DPWA Multichannel	

### $\Delta(1910)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $\Delta(1910) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS (GeV <sup>−1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.027±0.009	−30 ± 60	SOKHOYAN	15A	DPWA Multichannel

### $\Delta(1910)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $\Delta(1910) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>−1/2</sup> )	DOCUMENT ID	TECN	COMMENT
<b>+0.020±0.010 OUR ESTIMATE</b>			
0.026±0.008	SOKHOYAN	15A	DPWA Multichannel
−0.002±0.008	ARNDT	96	IPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.026±0.008	GUTZ	14	DPWA Multichannel
0.022±0.009	ANISOVICH	12A	DPWA Multichannel
0.030±0.002	SHRESTHA	12A	DPWA Multichannel

### $\Delta(1910)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $\Delta(1910)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	(KSU)
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(GWU)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(PITT, ANL)
VRANA	00	PR D20 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(VPI)
ARNDT	96	PR C53 430	R.A. Arndt, I.I. Strakovsky, R.L. Workman	(KARL)
HOEHLER	79	$\pi N$ Newsletter 9 1	G. Hohler	(CMU, LBL) IJP
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(KARLT) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	

## $\Delta(1920)$ 3/2<sup>+</sup>

$$I(J^P) = \frac{3}{2}(\frac{3}{2}^+) \text{ Status: } ** *$$

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $\Delta(1920)$ POLE POSITION

REAL PART	DOCUMENT ID	TECN	COMMENT
<b>VALUE (MeV)</b>			
<b>1850 to 1950 (≈ 1900) OUR ESTIMATE</b>			
1875 ± 30	SOKHOYAN	15A	DPWA Multichannel
1906 ± 10 ± 2	SVARC	14	L+P $\pi N \rightarrow \pi N$
1900	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1900 ± 80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1875 ± 30	GUTZ	14	DPWA Multichannel
1890 ± 30	ANISOVICH	12A	DPWA Multichannel
2110	SHRESTHA	12A	DPWA Multichannel
1880	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART	DOCUMENT ID	TECN	COMMENT
<b>VALUE (MeV)</b>			
<b>200 to 400 (≈ 300) OUR ESTIMATE</b>			
300 ± 40	SOKHOYAN	15A	DPWA Multichannel
310 ± 20 ± 11	SVARC	14	L+P $\pi N \rightarrow \pi N$
300±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
300 ± 40	GUTZ	14	DPWA Multichannel
300 ± 60	ANISOVICH	12A	DPWA Multichannel
386	SHRESTHA	12A	DPWA Multichannel
120	VRANA	00	DPWA Multichannel

See key on page 601

Baryon Particle Listings  
 $\Delta(1920)$

$\Delta(1920)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
16±6	SOKHOYAN	15A	DPWA Multichannel
26±3±2	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
24±4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
16±6	GUTZ	14	DPWA Multichannel
17±8	ANISOVICH	12A	DPWA Multichannel

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−50±25	SOKHOYAN	15A	DPWA Multichannel
−130±5±3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−150±30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−50±25	GUTZ	14	DPWA Multichannel
−40±20	ANISOVICH	12A	DPWA Multichannel

$\Delta(1920)$  INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow \Delta\eta$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.15±0.04	70±20	GUTZ	14	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.17±0.08	70±20	ANISOVICH	12A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow \Sigma K$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.09±0.03	80±40	ANISOVICH	12A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow \Delta\pi, P\text{-wave}$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.20±0.08	−105±25	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.20±0.12	−120±30	ANISOVICH	12A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow \Delta\pi, F\text{-wave}$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.37±0.10	−90±20	SOKHOYAN	15A	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.28±0.07	−95±35	ANISOVICH	12A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow N(1535)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.03±0.02	35±45	GUTZ	14	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow N_{a_0}(980)$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.03±0.02	−85±45	GUTZ	14	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow N(1440)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.04±0.03	undefined	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1920) \rightarrow N(1520)\pi, S\text{-wave}$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.05±0.05	undefined	SOKHOYAN	15A	DPWA Multichannel

$\Delta(1920)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1900 to 1970 (<math>\approx 1920</math>) OUR ESTIMATE</b>			
1880±30	SOKHOYAN	15A	DPWA Multichannel
1920±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1868±10	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1880±30	GUTZ	14	DPWA Multichannel
1900±30	ANISOVICH	12A	DPWA Multichannel
2146±32	SHRESTHA	12A	DPWA Multichannel
2057±1	PENNER	02c	DPWA Multichannel
1889±100	VRANA	00	DPWA Multichannel

$\Delta(1920)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>180 to 300 (<math>\approx 260</math>) OUR ESTIMATE</b>			
300±40	SOKHOYAN	15A	DPWA Multichannel
300±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
220±80	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

300±40	GUTZ	14	DPWA Multichannel
310±60	ANISOVICH	12A	DPWA Multichannel
400±80	SHRESTHA	12A	DPWA Multichannel
525±32	PENNER	02c	DPWA Multichannel
123±53	VRANA	00	DPWA Multichannel

$\Delta(1920)$  DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	5–20 %
$\Gamma_2$ $\Sigma K$	2–6 %
$\Gamma_3$ $N\pi\pi$	
$\Gamma_4$ $\Delta(1232)\pi$	50–90 %
$\Gamma_5$ $\Delta(1232)\pi, P\text{-wave}$	8–28 %
$\Gamma_6$ $\Delta(1232)\pi, F\text{-wave}$	44–72 %
$\Gamma_7$ $N(1440)\pi, P\text{-wave}$	<4 %
$\Gamma_8$ $N(1520)\pi, S\text{-wave}$	<5 %
$\Gamma_9$ $N(1535)\pi$	<2 %
$\Gamma_{10}$ $N_{a_0}(980)$	seen
$\Gamma_{11}$ $\Delta(1232)\eta$	5–17 %

$\Delta(1920)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
<b>5 to 20 OUR ESTIMATE</b>				
8±4	SOKHOYAN	15A	DPWA Multichannel	
20±5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
14±4	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
8±4	GUTZ	14	DPWA Multichannel	
8±4	ANISOVICH	12A	DPWA Multichannel	
16±4	SHRESTHA	12A	DPWA Multichannel	
15±1	PENNER	02c	DPWA Multichannel	
5±4	VRANA	00	DPWA Multichannel	

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_2/\Gamma$
VALUE (%)				
4±2	ANISOVICH	12A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.1±0.3	PENNER	02c	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
VALUE (%)				
18±10	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
22±12	ANISOVICH	12A	DPWA Multichannel	
7±5	SHRESTHA	12A	DPWA Multichannel	
41±3	VRANA	00	DPWA Multichannel	

$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_6/\Gamma$
VALUE (%)				
58±14	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
45±20	ANISOVICH	12A	DPWA Multichannel	

$\Gamma(N(1440)\pi, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE (%)				
<4	SOKHOYAN	15A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<20	SHRESTHA	12A	DPWA Multichannel	
53±8	VRANA	00	DPWA Multichannel	

$\Gamma(N(1520)\pi, S\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE (%)				
<5	SOKHOYAN	15A	DPWA Multichannel	

$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_9/\Gamma$
VALUE (%)				
<2	GUTZ	14	DPWA Multichannel	

$\Gamma(N_{a_0}(980))/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_{10}/\Gamma$
VALUE (%)				
4±2	HORN	08A	DPWA Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •				

# Baryon Particle Listings

## $\Delta(1920)$ , $\Delta(1930)$

$\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}$					$\Gamma_{11}/\Gamma$
VALUE (%)	DOCUMENT	ID	TECN	COMMENT	
11 ± 6	GUTZ	14	DPWA	Multichannel	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
15 ± 8	ANISOVICH	12a	DPWA	Multichannel	

### $\Delta(1920)$ PHOTON DECAY AMPLITUDES AT THE POLE

#### $\Delta(1920) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT	
0.110±0.030	−50 ± 20	SOKHOYAN	15A	DPWA	Multichannel

#### $\Delta(1920) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT	
−0.100±0.040	0 ± 20	SOKHOYAN	15A	DPWA	Multichannel

### $\Delta(1920)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $\Delta(1920) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
0.110±0.030	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.110±0.030	GUTZ	14	DPWA	Multichannel
0.130 <sup>+0.030</sup> <sub>−0.060</sub>	ANISOVICH	12A	DPWA	Multichannel
0.051±0.010	SHRESTHA	12A	DPWA	Multichannel
−0.007	PENNER	02D	DPWA	Multichannel

#### $\Delta(1920) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
−0.105±0.035	SOKHOYAN	15A	DPWA	Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.105±0.035	GUTZ	14	DPWA	Multichannel
−0.115 <sup>+0.025</sup> <sub>−0.050</sub>	ANISOVICH	12A	DPWA	Multichannel
0.017±0.015	SHRESTHA	12A	DPWA	Multichannel
−0.001	PENNER	02D	DPWA	Multichannel

### $\Delta(1920)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $\Delta(1920)$ REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
HORN	08A	EPJ A38 173	I. Horn <i>et al.</i>	(CB-ELSA Collab.)
Also		PRL 101 202002	I. Horn <i>et al.</i>	(CB-ELSA Collab.)
PENNER	02C	PR C66 055211	G. Penner, U. Mosel	(GIES)
PENNER	02D	PR C66 055212	G. Penner, U. Mosel	(GIES)
VRANA	00	PRPL 328 161	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IUP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IUP
Also		Toronto Conf. 3	R. Koch	(KARLT) IUP

## $\Delta(1930)$ 5/2<sup>−</sup>

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $\Delta(1930)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1840 to 1960 (≈ 1900) OUR ESTIMATE			
1848± 9±19	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2001	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1850	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
1890±50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1882	SHRESTHA	12A	DPWA    Multichannel
1883	VRANA	00	DPWA    Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
175 to 360 (≈ 270) OUR ESTIMATE			
321±17±7	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
387	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
180	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
260±60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
187	SHRESTHA	12A	DPWA    Multichannel
250	VRANA	00	DPWA    Multichannel

### $\Delta(1930)$ ELASTIC POLE RESIDUE

#### MODULUS $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>8 to 20 (<math>\approx</math> 14) OUR ESTIMATE</b>			
$9 \pm 1$	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
7	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
20	HOEHLER	93	SPED $\pi N \rightarrow \pi N$
$18 \pm 6$	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

#### PHASE $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−10 to −40 (≈ − 30) OUR ESTIMATE			
−37± 3±7	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−12	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−20±40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

### $\Delta(1930)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1900 to 2000 (≈ 1950) OUR ESTIMATE			
2233 ± 53	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1940 ± 30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1901 ± 15	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1930 ± 12	SHRESTHA	12A	DPWA Multichannel
1932 ± 100	VRANA	00	DPWA Multichannel

### $\Delta(1930)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
220 to 500 (≈ 360) OUR ESTIMATE			
773±187	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
320± 60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
195± 60	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
235± 39	SHRESTHA	12A	DPWA Multichannel
316±237	VRANA	00	DPWA Multichannel

### $\Delta(1930)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	5–15 %
$\Gamma_2$ $N\gamma$	0.0–0.01 %
$\Gamma_3$ $N\gamma$ , helicity=1/2	0.0–0.005 %
$\Gamma_4$ $N\gamma$ , helicity=3/2	0.0–0.004 %

### $\Delta(1930)$ BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>5 to 15 OUR ESTIMATE</b>				
8.1±1.2	ARNDT	06	DPWA	$\pi N \rightarrow \pi N, \eta N$
14 ±4	CUTKOSKY	80	IPWA	$\pi N \rightarrow \pi N$
4 ±3	HOEHLER	79	IPWA	$\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7.9±0.4	SHRESTHA	12a	DPWA	Multichannel
9 ±8	VRANA	00	DPWA	Multichannel

### $\Delta(1930)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

#### $\Delta(1930) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT	
−0.007±0.010	ARNDT	96	IPWA	$\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.011±0.003	SHRESTHA	12A	DPWA	Multichannel

See key on page 601

Baryon Particle Listings  
 $\Delta(1930)$ ,  $\Delta(1940)$

$\Delta(1930) \rightarrow N\gamma$ , helicity-3/2 amplitude  $A_{3/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.005±0.010	ARNDT	96	IPWA $\gamma N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.002±0.002	SHRESTHA	12A	DPWA Multichannel

$\Delta(1930)$  FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

$\Delta(1930)$  REFERENCES

For early references, see Physics Letters **111B** 1 (1982).

PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>	(GWU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
ARNDT	96	PR C53 430	R.A. Arndt, I.I. Strakovsky, R.L. Workman	(VPI)
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler	(KARL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(1940) 3/2^-$

$I(J^P) = \frac{3}{2}(\frac{3}{2}^-)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

$\Delta(1940)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2040±50	SOKHOYAN	15A	DPWA Multichannel
1878±11±5.5	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1900±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2040±50	GUTZ	14	DPWA Multichannel
1990 <sup>+100</sup> <sub>-50</sub>	ANISOVICH	12A	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
450±90	SOKHOYAN	15A	DPWA Multichannel
212±21±6	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
200±60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
450±90	GUTZ	14	DPWA Multichannel
450±90	ANISOVICH	12A	DPWA Multichannel

$\Delta(1940)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
6±3	SOKHOYAN	15A	DPWA Multichannel
9±1±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
8±3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
4±3	GUTZ	14	DPWA Multichannel
4±4	ANISOVICH	12A	DPWA Multichannel

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−90±35	SOKHOYAN	15A	DPWA Multichannel
140±7±7	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
135±45	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−50±35	GUTZ	14	DPWA Multichannel

$\Delta(1940)$  INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

Normalized residue in  $N\pi \rightarrow \Delta(1940) \rightarrow \Delta(1232)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
<0.01	undefined	GUTZ	14	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1940) \rightarrow N(1535)\pi$

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
<0.03	undefined	GUTZ	14	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1940) \rightarrow \Delta(1232)\pi$ , S-wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.12±0.06	120±45	SOKHOYAN	15A	DPWA Multichannel

Normalized residue in  $N\pi \rightarrow \Delta(1940) \rightarrow \Delta(1232)\pi$ , D-wave

MODULUS	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.06±0.04	−80±35	SOKHOYAN	15A	DPWA Multichannel

$\Delta(1940)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1940 to 2060 (≈ 2000) OUR ESTIMATE</b>			
2050±40	SOKHOYAN	15A	DPWA Multichannel
1940±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2050±40	GUTZ	14	DPWA Multichannel
1995 <sup>+105</sup> <sub>-60</sub>	ANISOVICH	12A	DPWA Multichannel

$\Delta(1940)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
450±70	SOKHOYAN	15A	DPWA Multichannel
200±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
450±70	GUTZ	14	DPWA Multichannel
450±100	ANISOVICH	12A	DPWA Multichannel

$\Delta(1940)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	1–7 %
$\Gamma_2$ $N\pi\pi$	
$\Gamma_3$ $\Delta(1232)\pi$	30–85 %
$\Gamma_4$ $\Delta(1232)\pi$ , S-wave	25–65 %
$\Gamma_5$ $\Delta(1232)\pi$ , D-wave	5–20 %
$\Gamma_6$ $N(1535)\pi$	2–14 %
$\Gamma_7$ $N_{a_0}(980)$	seen
$\Gamma_8$ $\Delta(1232)\eta$	4–16 %
$\Gamma_9$ $N\gamma$ , helicity=1/2	seen
$\Gamma_{10}$ $N\gamma$ , helicity=3/2	seen

$\Delta(1940)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
2±1	SOKHOYAN	15A	DPWA Multichannel
5±2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2±1	GUTZ	14	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, S\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
46±20	SOKHOYAN	15A	DPWA Multichannel

$\Gamma(\Delta(1232)\pi, D\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
12±7	SOKHOYAN	15A	DPWA Multichannel

$\Gamma(N(1535)\pi)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
8±6	GUTZ	14	DPWA Multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2±1	HORN	08A	DPWA Multichannel

$\Gamma(N_{a_0}(980))/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2±1	HORN	08A	DPWA Multichannel

$\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
10±6	GUTZ	14	DPWA Multichannel
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
4±2	HORN	08A	DPWA Multichannel

$\Delta(1940)$  PHOTON DECAY AMPLITUDES AT THE POLE

$\Delta(1940) \rightarrow N\gamma$ , helicity-1/2 amplitude  $A_{1/2}$

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.170±0.120 −0.100	−10±30	SOKHOYAN	15A	DPWA Multichannel

# Baryon Particle Listings

## $\Delta(1940)$ , $\Delta(1950)$

### $\Delta(1940) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$

MODULUS (GeV <sup>-1/2</sup> )	PHASE (°)	DOCUMENT ID	TECN	COMMENT
0.150±0.080	−10 ± 30	SOKHOYAN	15A	DPWA Multichannel

### $\Delta(1940)$ BREIT-WIGNER PHOTON DECAY AMPLITUDES

### $\Delta(1940) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.170 <sup>+0.110</sup> <sub>−0.080</sub>	SOKHOYAN	15A	DPWA Multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.170 <sup>+0.110</sup> <sub>−0.080</sub>	GUTZ	14	DPWA Multichannel
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### $\Delta(1940) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
0.150±0.080	SOKHOYAN	15A	DPWA Multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.150±0.080	GUTZ	14	DPWA Multichannel
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### $\Delta(1940)$ FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

### $\Delta(1940)$ REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan <i>et al.</i>	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz <i>et al.</i>	(CBELSA/TAPS Collab.)
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
HORN	08A	EPJ A38 173	I. Horn <i>et al.</i>	(CB-ELSA Collab.)
Also		PRL 101 202002	I. Horn <i>et al.</i>	(CB-ELSA Collab.)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohlner <i>et al.</i>	(KARLT)

## $\Delta(1950)$ 7/2<sup>+</sup>

$I(J^P) = \frac{3}{2}(\frac{7}{2}^+)$  Status: \* \* \* \*

Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C **38** 070001 (2014).

### $\Delta(1950)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 1870 to 1890 (≈ 1880) OUR ESTIMATE

1888 ± 4	SOKHOYAN	15A	DPWA Multichannel
1877 ± 2±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
1876	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1878	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
1890 ±15	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1888 ± 4	GUTZ	14	DPWA Multichannel
1890 ± 4	ANISOVICH	12A	DPWA Multichannel
1871	SHRESTHA	12A	DPWA Multichannel
1910	VRANA	00	DPWA Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 220 to 260 (≈ 240) OUR ESTIMATE

245 ± 8	SOKHOYAN	15A	DPWA Multichannel
223 ± 4±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
227	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
230	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
260±40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

245 ± 8	GUTZ	14	DPWA Multichannel
243 ± 8	ANISOVICH	12A	DPWA Multichannel
220	SHRESTHA	12A	DPWA Multichannel
230	VRANA	00	DPWA Multichannel

### $\Delta(1950)$ ELASTIC POLE RESIDUE

#### MODULUS $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 44 to 60 (≈ 52) OUR ESTIMATE

58±2	SOKHOYAN	15A	DPWA Multichannel
44±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
53	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
47	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
50±7	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

58±2	GUTZ	14	DPWA Multichannel
58±2	ANISOVICH	12A	DPWA Multichannel

### PHASE $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
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#### −24 to −40 (≈ − 32) OUR ESTIMATE

−24 ± 3	SOKHOYAN	15A	DPWA Multichannel
−39 ±1 ±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−31	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−32	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
−33 ± 8	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

−24 ± 3	GUTZ	14	DPWA Multichannel
−24 ± 3	ANISOVICH	12A	DPWA Multichannel

### $\Delta(1950)$ INELASTIC POLE RESIDUE

The “normalized residue” is the residue divided by  $\Gamma_{pole}/2$ .

### Normalized residue in $N\pi \rightarrow \Delta(1950) \rightarrow \Sigma K$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
5 ± 1	− 65 ± 25	ANISOVICH	12A	DPWA Multichannel

### Normalized residue in $N\pi \rightarrow \Delta(1950) \rightarrow \Delta\pi$ , $F$ -wave

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
12 ± 4	undefined	SOKHOYAN	15A	DPWA Multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •

12 ± 4	12 ± 10	ANISOVICH	12A	DPWA Multichannel
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### Normalized residue in $N\pi \rightarrow \Delta(1950) \rightarrow \Delta(1232)\eta$

MODULUS (%)	PHASE (°)	DOCUMENT ID	TECN	COMMENT
3.5 ± 0.5	90 ± 25	GUTZ	14	DPWA Multichannel

### $\Delta(1950)$ BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 1915 to 1950 (≈ 1930) OUR ESTIMATE

1917 ± 4	SOKHOYAN	15A	DPWA Multichannel
1921.3 ± 0.2	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
1950 ± 15	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
1913 ± 8	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1917 ± 4	GUTZ	14	DPWA Multichannel
1915 ± 6	ANISOVICH	12A	DPWA Multichannel
1918 ± 1	SHRESTHA	12A	DPWA Multichannel
1936 ± 5	VRANA	00	DPWA Multichannel

### $\Delta(1950)$ BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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#### 235 to 335 (≈ 285) OUR ESTIMATE

251 ± 8	SOKHOYAN	15A	DPWA Multichannel
271.1 ± 1.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
340 ± 50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
224 ± 10	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

251 ± 8	GUTZ	14	DPWA Multichannel
246 ± 10	ANISOVICH	12A	DPWA Multichannel
259 ± 4	SHRESTHA	12A	DPWA Multichannel
245 ± 12	VRANA	00	DPWA Multichannel

### $\Delta(1950)$ DECAY MODES

The following branching fractions are our estimates, not fits or averages.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	35–45 %
$\Gamma_2$ $\Sigma K$	0.3–0.5 %
$\Gamma_3$ $N\pi\pi$	
$\Gamma_4$ $\Delta(1232)\pi$ , $F$ -wave	1–9 %
$\Gamma_5$ $N(1680)\pi$ , $P$ -wave	3–9 %
$\Gamma_6$ $\Delta(1232)\eta$	< 1 %

### $\Delta(1950)$ BRANCHING RATIOS

#### $\Gamma(N\pi)/\Gamma_{\text{total}}$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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#### 35 to 45 OUR ESTIMATE

46 ± 2	SOKHOYAN	15A	DPWA Multichannel
47.1 ± 0.1	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
39 ± 4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
38 ± 2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

• • • We do not use the following data for averages, fits, limits, etc. • • •

46 ± 2	GUTZ	14	DPWA Multichannel
45 ± 2	ANISOVICH	12A	DPWA Multichannel
45.6 ± 0.4	SHRESTHA	12A	DPWA Multichannel
44 ± 1	VRANA	00	DPWA Multichannel

#### $\Gamma_1/\Gamma$



See key on page 601

Baryon Particle Listings  
 $\Delta(1950)$ ,  $\Delta(2000)$

$\Gamma(\Sigma K)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
0.4 ± 0.1	ANISOVICH	12A	DPWA Multichannel				
$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
5 ± 4	SOKHOYAN	15A	DPWA Multichannel				
• • • We do not use the following data for averages, fits, limits, etc. • • •							
2.8 ± 1.4	ANISOVICH	12A	DPWA Multichannel				
8 ± 1	SHRESTHA	12A	DPWA Multichannel				
36 ± 1	VRANA	00	DPWA Multichannel				
$\Gamma(N(1680)\pi, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
6 ± 3	SOKHOYAN	15A	DPWA Multichannel				
$\Gamma(\Delta(1232)\eta)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
<1	GUTZ	14	DPWA Multichannel				

$\Delta(1950)$  PHOTON DECAY AMPLITUDES AT THE POLE

$\Delta(1950) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
−0.067 ± 0.004	−10 ± 5	SOKHOYAN	15A	DPWA Multichannel
$\Delta(1950) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
MODULUS ( $\text{GeV}^{-1/2}$ )	PHASE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
−0.095 ± 0.004	−10 ± 5	SOKHOYAN	15A	DPWA Multichannel

$\Delta(1950)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES

$\Delta(1950) \rightarrow N\gamma$ , helicity-1/2 amplitude $A_{1/2}$				
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT	
−0.067 ± 0.005	SOKHOYAN	15A	DPWA Multichannel	
−0.083 ± 0.004	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.067 ± 0.005	GUTZ	14	DPWA Multichannel	
−0.071 ± 0.004	ANISOVICH	12A	DPWA Multichannel	
−0.065 ± 0.001	SHRESTHA	12A	DPWA Multichannel	
−0.094	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$	
$\Delta(1950) \rightarrow N\gamma$ , helicity-3/2 amplitude $A_{3/2}$				
VALUE ( $\text{GeV}^{-1/2}$ )	DOCUMENT ID	TECN	COMMENT	
−0.094 ± 0.004	SOKHOYAN	15A	DPWA Multichannel	
−0.096 ± 0.004	WORKMAN	12A	DPWA $\gamma N \rightarrow N\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.094 ± 0.004	GUTZ	14	DPWA Multichannel	
−0.094 ± 0.005	ANISOVICH	12A	DPWA Multichannel	
−0.083 ± 0.001	SHRESTHA	12A	DPWA Multichannel	
−0.121	DRECHSEL	07	DPWA $\gamma N \rightarrow \pi N$	

$\Delta(1950)$  FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

$\Delta(1950)$  REFERENCES

SOKHOYAN	15A	EPJ A51 95	V. Sokhoyan et al.	(CBELSA/TAPS Collab.)
GUTZ	14	EPJ A50 74	E. Gutz et al.	(CBELSA/TAPS Collab.)
PDG	14	CPC 38 070001	K. Olive et al.	(PDG Collab.)
SVARC	14	PR C89 045205	A. Svarc et al.	
ANISOVICH	12A	EPJ A48 15	A.V. Anisovich et al.	(BONN, PNPI)
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	(KSU)
WORKMAN	12A	PR C86 015202	R. Workman et al.	(GWU)
DRECHSEL	07	EPJ A34 69	D. Drechsel, S.S. Kamalov, L. Tiator	(MAINZ, JINR)
ARNDT	06	PR C74 045205	R.A. Arndt et al.	(GWU)
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
HOEHLER	79	$\pi N$ Newsletter 9 1	G. Hohlner et al.	(KARLT) IJP
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky et al.	(CMU, LBL) IJP
Also		PR D30 2839	R.E. Cutkosky et al.	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohlner et al.	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(2000) 5/2^+$

$I(J^P) = \frac{3}{2}(\frac{5}{2}^+)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

$\Delta(2000)$  POLE POSITION

REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1998 ± 4 ± 4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2150 ± 100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1976	SHRESTHA	12A	DPWA Multichannel
1697	VRANA	00	DPWA Multichannel

−2×IMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
404 ± 10 ± 4	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
350 ± 100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
488	SHRESTHA	12A	DPWA Multichannel
112	VRANA	00	DPWA Multichannel

$\Delta(2000)$  ELASTIC POLE RESIDUE

MODULUS $ r $			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
34 ± 1 ± 1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
16 ± 5	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
PHASE $\theta$			
VALUE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
110 ± 1 ± 3	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
150 ± 90	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2000)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2200 ± 125	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2015 ± 24	SHRESTHA	12A	DPWA Multichannel
1724 ± 61	VRANA	00	DPWA Multichannel
1752 ± 32	MANLEY	92	IPWA $\pi N \rightarrow \pi N$ & $N\pi\pi$

$\Delta(2000)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
400 ± 125	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
500 ± 52	SHRESTHA	12A	DPWA Multichannel
138 ± 68	VRANA	00	DPWA Multichannel
251 ± 93	MANLEY	92	IPWA $\pi N \rightarrow \pi N$ & $N\pi\pi$

$\Delta(2000)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	3–11 %
$\Gamma_2$ $N\pi\pi$	
$\Gamma_3$ $\Delta(1232)\pi$ , $P$ -wave	seen
$\Gamma_4$ $\Delta(1232)\pi$ , $F$ -wave	seen
$\Gamma_5$ $N\rho$ , $S=3/2$ , $P$ -wave	seen
$\Gamma_6$ $N\gamma$	
$\Gamma_7$ $N\gamma$ , helicity=1/2	seen
$\Gamma_8$ $N\gamma$ , helicity=3/2	seen

$\Delta(2000)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
7 ± 4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$				
• • • We do not use the following data for averages, fits, limits, etc. • • •							
7 ± 1	SHRESTHA	12A	DPWA Multichannel				
0 ± 1	VRANA	00	DPWA Multichannel				
2 ± 1	MANLEY	92	IPWA $\pi N \rightarrow \pi N$ & $N\pi\pi$				

$\Gamma(\Delta(1232)\pi, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •							
3 ± 3	SHRESTHA	12A	DPWA Multichannel				
0 ± 1	VRANA	00	DPWA Multichannel				

$\Gamma(\Delta(1232)\pi, F\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •							
< 3	SHRESTHA	12A	DPWA Multichannel				
40 ± 1	VRANA	00	DPWA Multichannel				

$\Gamma(N\rho, S=3/2, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	VALUE (%)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •							
90 ± 3	SHRESTHA	12A	DPWA Multichannel				
60 ± 60	VRANA	00	DPWA Multichannel				

Baryon Particle Listings

$\Delta(2000)$ ,  $\Delta(2150)$ ,  $\Delta(2200)$ ,  $\Delta(2300)$

$\Delta(2000)$  BREIT-WIGNER PHOTON DECAY AMPLITUDES

$\Delta(2000) \rightarrow p \gamma$ , helicity-1/2 amplitude  $A_{1/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−0.061±0.018	SHRESTHA	12A	DPWA Multichannel

$\Delta(2000) \rightarrow p \gamma$ , helicity-3/2 amplitude  $A_{3/2}$

VALUE (GeV <sup>-1/2</sup> )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.158±0.032	SHRESTHA	12A	DPWA Multichannel

$\Delta(2000)$  FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

$\Delta(2000)$  REFERENCES

SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	
SHRESTHA	12A	PR C86 055203	M. Shrestha, D.M. Manley	
VRANA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
MANLEY	92	PR D45 4002	D.M. Manley, E.M. Saleski	(KSA) IJP
Also		PR D30 904	D.M. Manley <i>et al.</i>	(VPI)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT)

$\Delta(2150)$  1/2<sup>−</sup>

$I(J^P) = \frac{3}{2}(\frac{1}{2}^-)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Delta(2150)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2140±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
200±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2150)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
7±2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−60±90	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2150)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2150±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2150)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
200±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2150)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N \pi$	6–10 %

$\Delta(2150)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
8±2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	

$\Delta(2150)$  REFERENCES

CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)

$\Delta(2200)$  7/2<sup>−</sup>

$I(J^P) = \frac{3}{2}(\frac{7}{2}^-)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Delta(2200)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2100±50	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
340±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2200)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
8±3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−70±40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2200)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2200±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2215±60	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2200)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
450±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
400±100	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2200)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N \pi$	3–8 %

$\Delta(2200)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
6±2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
5±2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	

$\Delta(2200)$  REFERENCES

CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(2300)$  9/2<sup>+</sup>

$I(J^P) = \frac{3}{2}(\frac{9}{2}^+)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

$\Delta(2300)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2370±80	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
420±160	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2300)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
10±4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

See key on page 601

Baryon Particle Listings  
 $\Delta(2300)$ ,  $\Delta(2350)$ ,  $\Delta(2390)$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−20±30	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2300)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2400±125	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2217±80	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2300)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
425±150	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
300±100	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2300)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	1–8 %

$\Delta(2300)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
6±2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
3±2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	

$\Delta(2300)$  REFERENCES

CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(2350)$   $5/2^-$

$I(J^P) = \frac{3}{2}(\frac{5}{2}^-)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Delta(2350)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2400±125	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2427	VRA NA	00	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
400±150	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
458	VRA NA	00	DPWA Multichannel

$\Delta(2350)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
15±8	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−70±70	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2350)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2400±125	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2305±26	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2459±100	VRA NA	00	DPWA Multichannel

$\Delta(2350)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
400±150	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
300±70	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
480±360	VRA NA	00	DPWA Multichannel

$\Delta(2350)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	4–30 %

$\Delta(2350)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
20±10	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
4±2	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
7±14	VRA NA	00	DPWA Multichannel	

$\Delta(2350)$  REFERENCES

VRA NA	00	PRPL 328 181	T.P. Vrana, S.A. Dytman, T.-S.H. Lee	(PITT, ANL)
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL) IJP
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(2390)$   $7/2^+$

$I(J^P) = \frac{3}{2}(\frac{7}{2}^+)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Delta(2390)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2223±15±19	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2350±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
431±26±7	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
260±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2390)$  ELASTIC POLE RESIDUE

MODULUS  $|r|$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
26±2±1	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
12±6	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

PHASE  $\theta$

VALUE (°)	DOCUMENT ID	TECN	COMMENT
−160±5±11	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−90±60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2390)$  BREIT-WIGNER MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2350±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2425±60	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2390)$  BREIT-WIGNER WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
300±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
300±80	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2390)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	3–12 %

$\Delta(2390)$  BRANCHING RATIOS

$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE (%)				
8±4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$	
7±4	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$	

$\Delta(2390)$  FOOTNOTES

<sup>1</sup> Fit to the amplitudes of HOEHLER 79.

$\Delta(2390)$  REFERENCES

SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>	(CMU, LBL) IJP
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>	(CMU, LBL)
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

Baryon Particle Listings

$\Delta(2400)$ ,  $\Delta(2420)$ ,  $\Delta(2750)$

<div><math>\Delta(2400)</math> <math>9/2^-</math></div>	$I(J^P) = \frac{3}{2}(\frac{9}{2}^-)$ Status: **
OMITTED FROM SUMMARY TABLE	

$\Delta(2400)$ POLE POSITION			
REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1983	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2260±60	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
−2×IMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
878	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
320±160	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2400)$ ELASTIC POLE RESIDUE			
MODULUS $ r $			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
24	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
8±4	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
PHASE $\theta$			
VALUE (°)	DOCUMENT ID	TECN	COMMENT
−139	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−25±15	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2400)$ BREIT-WIGNER MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2643±141	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2300±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2468±50	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2400)$ BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
895±432	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
330±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
480±100	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2400)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	3–9 %

$\Delta(2400)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)			
6.4±2.2	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
5±2	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
6±3	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2400)$ REFERENCES			
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>
Also		Toronto Conf. 3	R. Koch
			(GWU)
			(CMU, LBL) IJP
			(CMU, LBL)
			(KARLT) IJP
			(KARLT) IJP

<div><math>\Delta(2420)</math> <math>11/2^+</math></div>	$I(J^P) = \frac{3}{2}(\frac{11}{2}^+)$ Status: ****
Older and obsolete values are listed and referenced in the 2014 edition, Chinese Physics C <b>38</b> 070001 (2014).	

$\Delta(2420)$ POLE POSITION			
REAL PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2260 to 2400 (≈ 2330) OUR ESTIMATE			
2454±4±11	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
2529	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2300	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
2360±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

−2×IMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
350 to 750 (≈ 550) OUR ESTIMATE			
462±8±50	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
621	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
620	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
420±100	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2420)$ ELASTIC POLE RESIDUE			
MODULUS $ r $			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
20 to 40 (≈ 30) OUR ESTIMATE			
30±1±7	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
33	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
39	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
18±6	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

PHASE $\theta$			
VALUE (°)	DOCUMENT ID	TECN	COMMENT
−60 to 20 (≈ −20) OUR ESTIMATE			
11±1±8	<sup>1</sup> SVARC	14	L+P $\pi N \rightarrow \pi N$
−45	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
−60	HOEHLER	93	ARGD $\pi N \rightarrow \pi N$
−30±40	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$

$\Delta(2420)$ BREIT-WIGNER MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
2300 to 2500 (≈ 2420) OUR ESTIMATE			
2633±29	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
2400±125	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
2416±17	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2420)$ BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
300 to 500 (≈ 400) OUR ESTIMATE			
692±47	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
450±150	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
340±28	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2420)$ DECAY MODES	
The following branching fractions are our estimates, not fits or averages.	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\pi$	5–15 %

$\Delta(2420)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE (%)			
5 to 15 OUR ESTIMATE			
8.5±0.8	ARNDT	06	DPWA $\pi N \rightarrow \pi N, \eta N$
8±3	CUTKOSKY	80	IPWA $\pi N \rightarrow \pi N$
8.0±1.5	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2420)$ FOOTNOTES	
<sup>1</sup> Fit to the amplitudes of HOEHLER 79.	

$\Delta(2420)$ REFERENCES			
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>
SVARC	14	PR C89 045205	A. Svarc <i>et al.</i>
ARNDT	06	PR C74 045205	R.A. Arndt <i>et al.</i>
HOEHLER	93	$\pi N$ Newsletter 9 1	G. Hohler
CUTKOSKY	80	Toronto Conf. 19	R.E. Cutkosky <i>et al.</i>
Also		PR D20 2839	R.E. Cutkosky <i>et al.</i>
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>
Also		Toronto Conf. 3	R. Koch
			(PDG Collab.)
			(GWU)
			(KARL)
			(CMU, LBL) IJP
			(CMU, LBL)
			(KARLT) IJP
			(KARLT) IJP

<div><math>\Delta(2750)</math> <math>13/2^-</math></div>	$I(J^P) = \frac{3}{2}(\frac{13}{2}^-)$ Status: **		
OMITTED FROM SUMMARY TABLE			
$\Delta(2750)$ BREIT-WIGNER MASS			
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
2794 ± 80	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

See key on page 601

Baryon Particle Listings  
 $\Delta(2750)$ ,  $\Delta(2950)$ ,  $\Delta(\sim 3000)$

$\Delta(2750)$ BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$350 \pm 100$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2750)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad N\pi$	2–6 %

$\Delta(2750)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$4.0 \pm 1.5$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2750)$ REFERENCES				
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(2950) 15/2^+$	$I(J^P) = \frac{3}{2}(\frac{15}{2}^+)$ Status: **
OMITTED FROM SUMMARY TABLE	

$\Delta(2950)$ BREIT-WIGNER MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$2990 \pm 100$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2950)$ BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$330 \pm 100$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2950)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad N\pi$	2–6 %

$\Delta(2950)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$4 \pm 2$	HOEHLER	79	IPWA $\pi N \rightarrow \pi N$

$\Delta(2950)$ REFERENCES				
HOEHLER	79	PDAT 12-1	G. Hohler <i>et al.</i>	(KARLT) IJP
Also		Toronto Conf. 3	R. Koch	(KARLT) IJP

$\Delta(\sim 3000)$  Region  
Partial-Wave Analyses

OMITTED FROM SUMMARY TABLE  
We list here miscellaneous high-mass candidates for isospin-3/2 resonances found in partial-wave analyses.  
  
Our 1982 edition also had a  $\Delta(2850)$  and a  $\Delta(3230)$ . The evidence for them was deduced from total cross-section and  $180^\circ$  elastic cross-section measurements. The  $\Delta(2850)$  has been resolved into the  $\Delta(2750) \, l_{3,13}$  and  $\Delta(2950) \, K_{3,15}$ . The  $\Delta(3230)$  is perhaps related to the  $K_{3,13}$  of HENDRY 78 and to the  $L_{3,17}$  of KOCH 80.

$\Delta(\sim 3000)$ BREIT-WIGNER MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
3300	<sup>1</sup> KOCH	80	IPWA $\pi N \rightarrow \pi N \, L_{3,17}$ wave
3500	<sup>1</sup> KOCH	80	IPWA $\pi N \rightarrow \pi N \, M_{3,19}$ wave
$2850 \pm 150$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, l_{3,11}$ wave
$3200 \pm 200$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, K_{3,13}$ wave
$3300 \pm 200$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, L_{3,17}$ wave
$3700 \pm 200$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, M_{3,19}$ wave
$4100 \pm 300$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, N_{3,21}$ wave

$\Delta(\sim 3000)$ BREIT-WIGNER WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$700 \pm 200$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, l_{3,11}$ wave
$1000 \pm 300$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, K_{3,13}$ wave
$1100 \pm 300$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, L_{3,17}$ wave
$1300 \pm 400$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, M_{3,19}$ wave
$1600 \pm 500$	HENDRY	78	MPWA $\pi N \rightarrow \pi N \, N_{3,21}$ wave

$\Delta(\sim 3000)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad N\pi$	seen

$\Delta(\sim 3000)$ BRANCHING RATIOS			
$\Gamma(N\pi)/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma$
VALUE (%)	DOCUMENT ID	TECN	COMMENT
$6\pm 2$	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $l_{3,11}$ wave
$5\pm 2$	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $K_{3,13}$ wave
$3\pm 1$	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $L_{3,17}$ wave
$3\pm 1$	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $M_{3,19}$ wave
$2\pm 1$	HENDRY	78	MPWA $\pi N \rightarrow \pi N$ $N_{3,21}$ wave

$\Delta(\sim 3000)$  FOOTNOTES  
<sup>1</sup>In addition, KOCH 80 reports some evidence for an  $S_{31} \, \Delta(2700)$  and a  $P_{33} \, \Delta(2800)$ .

$\Delta(\sim 3000)$ REFERENCES				
KOCH	80	Toronto Conf. 3	R. Koch	(KARLT) IJP
HENDRY	78	PRL 41 222	A.W. Hendry	(IND, LBL) IJP
Also		ANP 136 1	A.W. Hendry	(IND)

Baryon Particle Listings

$\Lambda$

$\Lambda$  BARYONS  
( $S = -1, I = 0$ )  
 $\Lambda^0 = uds$



$I(J^P) = 0(\frac{1}{2}^+)$  Status: \*\*\*\*

We have omitted some results that have been superseded by later experiments. See our earlier editions.

$\Lambda$  MASS

The fit uses  $\Lambda$ ,  $\Sigma^+$ ,  $\Sigma^0$ ,  $\Sigma^-$  mass and mass-difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1115.683±0.006 OUR FIT</b>				
<b>1115.683±0.006 OUR AVERAGE</b>				
1115.678±0.006±0.006	20k	HARTOUNI	94	SPEC $pp$ 27.5 GeV/c
1115.690±0.008±0.006	18k	<sup>1</sup> HARTOUNI	94	SPEC $pp$ 27.5 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1115.59 ±0.08	935	HYMAN	72	HEBC
1115.39 ±0.12	195	MAYEUR	67	EMUL
1115.6 ±0.4		LONDON	66	HBC
1115.65 ±0.07	488	<sup>2</sup> SCHMIDT	65	HBC
1115.44 ±0.12		<sup>3</sup> BHOWMIK	63	RVUE

- <sup>1</sup>We assume  $CPT$  invariance: this is the  $\bar{\Lambda}$  mass as measured by HARTOUNI 94. See below for the fractional mass difference, testing  $CPT$ .
- <sup>2</sup>The SCHMIDT 65 masses have been reevaluated using our April 1973 proton and  $K^\pm$  and  $\pi^\pm$  masses. P. Schmidt, private communication (1974).
- <sup>3</sup>The mass has been raised 35 keV to take into account a 46 keV increase in the proton mass and an 11 keV decrease in the  $\pi^\pm$  mass (note added Reviews of Modern Physics **39** 1 (1967)).

$(m_\Lambda - m_{\bar{\Lambda}}) / m_\Lambda$

A test of  $CPT$  invariance.

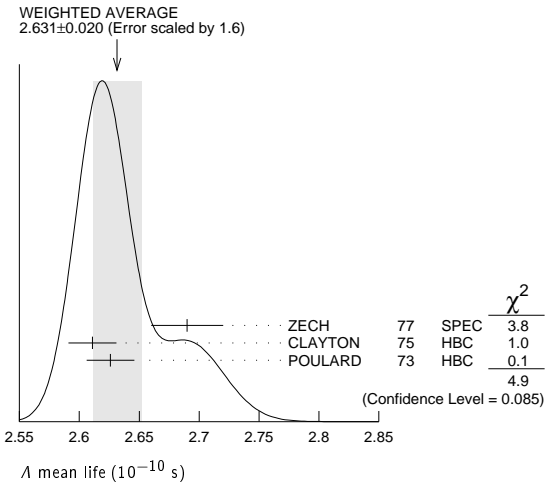
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>- 0.1 ± 1.1 OUR AVERAGE</b>				Error includes scale factor of 1.6.
- 1.3 ± 1.2	31k	<sup>1</sup> RYBICKI	96	NA32 $\pi^-$ Cu, 230 GeV
- 1.08 ± 0.90		HARTOUNI	94	SPEC $pp$ 27.5 GeV/c
4.5 ± 5.4		CHIEN	66	HBC 6.9 GeV/c $\bar{p}p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-26 ±13		BADIER	67	HBC 2.4 GeV/c $\bar{p}p$

<sup>1</sup>RYBICKI 96 is an analysis of old ACCMOR (NA32) data.

$\Lambda$  MEAN LIFE

Measurements with an error  $\geq 0.1 \times 10^{-10}$  s have been omitted altogether, and only the latest high-statistics measurements are used for the average.

VALUE ( $10^{-10}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.632±0.020 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
2.69 ±0.03	53k	ZECH	77	SPEC Neutral hyperon beam
2.611±0.020	34k	CLAYTON	75	HBC 0.96-1.4 GeV/c $K^- p$
2.626±0.020	36k	POULARD	73	HBC 0.4-2.3 GeV/c $K^- p$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.69 ±0.05	6582	ALTHOFF	73B	OSPK $\pi^+ n \rightarrow \Lambda K^+$
2.54 ±0.04	4572	BALTAY	71B	HBC $K^- p$ at rest
2.535±0.035	8342	GRIMM	68	HBC
2.47 ±0.08	2600	HEPP	68	HBC
2.35 ±0.09	916	BURAN	66	HLBC
2.452+0.056-0.054	2213	ENGELMANN	66	HBC
2.59 ±0.09	794	HUBBARD	64	HBC
2.59 ±0.07	1378	SCHWARTZ	64	HBC
2.36 ±0.06	2239	BLOCK	63	HEBC



$(\tau_\Lambda - \tau_{\bar{\Lambda}}) / \tau_\Lambda$

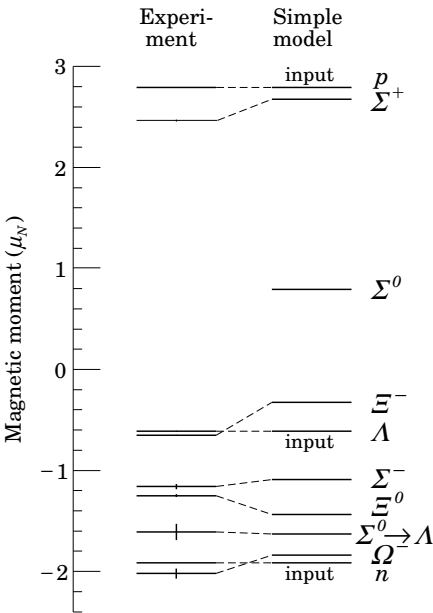
A test of  $CPT$  invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.001 ±0.009 OUR AVERAGE</b>			
-0.0018±0.0066±0.0056	BARNES	96	CNTR LEAR $\bar{p}p \rightarrow \bar{\Lambda}\Lambda$
0.044 ±0.085	BADIER	67	HBC 2.4 GeV/c $\bar{p}p$

BARYON MAGNETIC MOMENTS

Written 1994 by C.G. Wohl (LBNL).

The figure below shows the measured magnetic moments of the stable baryons. It also shows the predictions of the simplest



quark model, using the measured  $p$ ,  $n$ , and  $\Lambda$  moments as input. In this model, the moments are [1]

$$\begin{aligned} \mu_p &= (4\mu_u - \mu_d)/3 & \mu_n &= (4\mu_d - \mu_u)/3 \\ \mu_{\Sigma^+} &= (4\mu_u - \mu_s)/3 & \mu_{\Sigma^-} &= (4\mu_d - \mu_s)/3 \\ \mu_{\Sigma^0} &= (4\mu_s - \mu_u)/3 & \mu_{\Xi^-} &= (4\mu_s - \mu_d)/3 \\ \mu_\Lambda &= \mu_s & \mu_{\Sigma^0} &= (2\mu_u + 2\mu_d - \mu_s)/3 \\ \mu_{\Omega^-} &= 3\mu_s \end{aligned}$$

and the  $\Sigma^0 \rightarrow \Lambda$  transition moment is

$$\mu_{\Sigma^0\Lambda} = (\mu_d - \mu_u)/\sqrt{3}.$$

The quark moments that result from this model are  $\mu_u = +1.852 \mu_N$ ,  $\mu_d = -0.972 \mu_N$ , and  $\mu_s = -0.613 \mu_N$ . The corresponding effective quark masses, taking the quarks to be Dirac point particles, where  $\mu = q\hbar/2m$ , are 338, 322, and 510 MeV. As the figure shows, the model gives a good first approximation to the experimental moments. For efforts to make a better model, we refer to the literature [2].

## References

- See, for example, D.H. Perkins, *Introduction to High Energy Physics* (Addison-Wesley, Reading, MA, 1987), or D. Griffiths, *Introduction to Elementary Particles* (Harper & Row, New York, 1987).
- See, for example, J. Franklin, Phys. Rev. **D29**, 2648 (1984); H.J. Lipkin, Nucl. Phys. **B241**, 477 (1984); K. Suzuki, H. Kumagai, and Y. Tanaka, Europhys. Lett. **2**, 109 (1986); S.K. Gupta and S.B. Khadkikar, Phys. Rev. **D36**, 307 (1987); M.I. Krivoruchenko, Sov. J. Nucl. Phys. **45**, 109 (1987); L. Brekke and J.L. Rosner, Comm. Nucl. Part. Phys. **18**, 83 (1988); K.-T. Chao, Phys. Rev. **D41**, 920 (1990) and references cited therein. Also, see references cited in discussions of results in the experimental papers..

## Λ MAGNETIC MOMENT

See the "Note on Baryon Magnetic Moments" above. Measurements with an error  $\geq 0.15 \mu_N$  have been omitted.

VALUE ( $\mu_N$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.613 ± 0.004 OUR AVERAGE</b>				
-0.606 ± 0.015	200k	COX	81	SPEC
-0.6138 ± 0.0047	3M	SCHACHIN...	78	SPEC
-0.59 ± 0.07	350k	HELLER	77	SPEC
-0.57 ± 0.05	1.2M	BUNCE	76	SPEC
-0.66 ± 0.07	1300	DAHL-JENSEN	71	EMUL 200 kG field

## Λ ELECTRIC DIPOLE MOMENT

A nonzero value is forbidden by both  $T$  invariance and  $P$  invariance.

VALUE ( $10^{-16} e \text{ cm}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.5</b>	95	<sup>1</sup> PONDROM	81	SPEC
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<100	95	<sup>2</sup> BARONI	71	EMUL
<500	95	GIBSON	66	EMUL
<sup>1</sup> PONDROM 81 measures $(-3.0 \pm 7.4) \times 10^{-17} e\text{-cm}$ .				
<sup>2</sup> BARONI 71 measures $(-5.9 \pm 2.9) \times 10^{-15} e\text{-cm}$ .				

## Λ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \quad p\pi^-$	(63.9 ± 0.5) %	
$\Gamma_2 \quad n\pi^0$	(35.8 ± 0.5) %	
$\Gamma_3 \quad n\gamma$	( 1.75 ± 0.15) × 10 <sup>-3</sup>	
$\Gamma_4 \quad p\pi^-\gamma$	[a] ( 8.4 ± 1.4) × 10 <sup>-4</sup>	
$\Gamma_5 \quad p e^- \bar{\nu}_e$	( 8.32 ± 0.14) × 10 <sup>-4</sup>	
$\Gamma_6 \quad p\mu^- \bar{\nu}_\mu$	( 1.57 ± 0.35) × 10 <sup>-4</sup>	

### Lepton (L) and/or Baryon (B) number violating decay modes

$\Gamma_7 \quad \pi^+ e^-$	L,B	< 6	× 10 <sup>-7</sup>	90%
$\Gamma_8 \quad \pi^+ \mu^-$	L,B	< 6	× 10 <sup>-7</sup>	90%
$\Gamma_9 \quad \pi^- e^+$	L,B	< 4	× 10 <sup>-7</sup>	90%
$\Gamma_{10} \quad \pi^- \mu^+$	L,B	< 6	× 10 <sup>-7</sup>	90%
$\Gamma_{11} \quad K^+ e^-$	L,B	< 2	× 10 <sup>-6</sup>	90%

$\Gamma_{12} \quad K^+ \mu^-$	L,B	< 3	× 10 <sup>-6</sup>	90%
$\Gamma_{13} \quad K^- e^+$	L,B	< 2	× 10 <sup>-6</sup>	90%
$\Gamma_{14} \quad K^- \mu^+$	L,B	< 3	× 10 <sup>-6</sup>	90%
$\Gamma_{15} \quad K_S^0 \nu$	L,B	< 2	× 10 <sup>-5</sup>	90%
$\Gamma_{16} \quad \bar{p}\pi^+$	B	< 9	× 10 <sup>-7</sup>	90%

[a] See the Listings below for the pion momentum range used in this measurement.

## CONSTRAINED FIT INFORMATION

An overall fit to 5 branching ratios uses 20 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 10.5$  for 16 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-100			
$x_3$	-2	-1		
$x_5$	46	-46	-1	
$x_6$	0	0	0	0
	$x_1$	$x_2$	$x_3$	$x_5$

## Λ BRANCHING RATIOS

$\Gamma(p\pi^-)/\Gamma(N\pi)$				$\Gamma_1/(\Gamma_1+\Gamma_2)$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.641 ± 0.005 OUR FIT</b>				
<b>0.640 ± 0.005 OUR AVERAGE</b>				
0.646 ± 0.008	4572	BALTAY	71B	HBC $K^- p$ at rest
0.635 ± 0.007	6736	DOYLE	69	HBC $\pi^- p \rightarrow \Lambda K^0$
0.643 ± 0.016	903	HUMPHREY	62	HBC
0.624 ± 0.030		CRAWFORD	59B	HBC $\pi^- p \rightarrow \Lambda K^0$

$\Gamma(n\pi^0)/\Gamma(N\pi)$				$\Gamma_2/(\Gamma_1+\Gamma_2)$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.359 ± 0.005 OUR FIT</b>				
<b>0.310 ± 0.028 OUR AVERAGE</b>				
0.35 ± 0.05		BROWN	63	HLBC
0.291 ± 0.034	75	CHRETIEN	63	HLBC

$\Gamma(n\gamma)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.75 ± 0.15 OUR FIT</b>				
<b>1.75 ± 0.15</b>	1816	LARSON	93	SPEC $K^- p$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.78 ± 0.24 <sup>+0.14</sup> / <sub>-0.16</sub>	287	NOBLE	92	SPEC See LARSON 93

$\Gamma(n\gamma)/\Gamma(n\pi^0)$				$\Gamma_3/\Gamma_2$
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.86 ± 0.74 ± 0.57	24	BIAGI	86	SPEC SPS hyperon beam

$\Gamma(p\pi^-\gamma)/\Gamma(p\pi^-)$				$\Gamma_4/\Gamma_1$
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.32 ± 0.22</b>	72	BAGGETT	72C	HBC $\pi^- < 95 \text{ MeV}/c$

$\Gamma(p e^- \bar{\nu}_e)/\Gamma(p\pi^-)$				$\Gamma_5/\Gamma_1$
VALUE (units 10 <sup>-3</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.301 ± 0.019 OUR FIT</b>				
<b>1.301 ± 0.019 OUR AVERAGE</b>				
1.335 ± 0.056	7111	BOURQUIN	83	SPEC SPS hyperon beam
1.313 ± 0.024	10k	WISE	80	SPEC
1.23 ± 0.11	544	LINDQUIST	77	SPEC $\pi^- p \rightarrow K^0 \Lambda$
1.27 ± 0.07	1089	KATZ	73	HBC
1.31 ± 0.06	1078	ALTHOFF	71	OSPK
1.17 ± 0.13	86	<sup>1</sup> CANTER	71	HBC $K^- p$ at rest
1.20 ± 0.12	143	<sup>2</sup> MALONEY	69	HBC
1.17 ± 0.18	120	<sup>2</sup> BAGLIN	64	FBC $K^- \text{ freon } 1.45 \text{ GeV}/c$
1.23 ± 0.20	150	<sup>2</sup> ELY	63	FBC
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.32 ± 0.15	218	<sup>1</sup> LINDQUIST	71	OSPK See LINDQUIST 77

<sup>1</sup> Changed by us from  $\Gamma(p e^- \bar{\nu}_e)/\Gamma(N\pi)$  assuming the authors used  $\Gamma(p\pi^-)/\Gamma_{\text{total}} = 2/3$ .

<sup>2</sup> Changed by us from  $\Gamma(p e^- \bar{\nu}_e)/\Gamma(N\pi)$  because  $\Gamma(p e^- \nu)/\Gamma(p\pi^-)$  is the directly measured quantity.

# Baryon Particle Listings

$\Lambda$

$\Gamma(\rho\mu^-\bar{\nu}_\mu)/\Gamma(N\pi)$		$\Gamma_6/(\Gamma_1+\Gamma_2)$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>1.57\pm0.35</math> OUR FIT</b>					
<b><math>1.57\pm0.35</math> OUR AVERAGE</b>					
$1.4 \pm 0.5$	14	BAGGETT	72B	HBC	$K^-\rho$ at rest
$2.4 \pm 0.8$	9	CANTER	71B	HBC	$K^-\rho$ at rest
$1.3 \pm 0.7$	3	LIND	64	RVUE	
$1.5 \pm 1.2$	2	RONNE	64	FBC	

## Lepton (L) and/or Baryon (B) number violating decay modes

$\Gamma(\pi^+e^-)/\Gamma_{\text{total}}$		$\Gamma_7/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6 \times 10^{-7}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(\pi^+\mu^-)/\Gamma_{\text{total}}$		$\Gamma_8/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6 \times 10^{-7}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(\pi^-e^+)/\Gamma_{\text{total}}$		$\Gamma_9/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;4 \times 10^{-7}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(\pi^-\mu^+)/\Gamma_{\text{total}}$		$\Gamma_{10}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;6 \times 10^{-7}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(K^+e^-)/\Gamma_{\text{total}}$		$\Gamma_{11}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2 \times 10^{-6}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(K^+\mu^-)/\Gamma_{\text{total}}$		$\Gamma_{12}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;3 \times 10^{-6}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(K^-e^+)/\Gamma_{\text{total}}$		$\Gamma_{13}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2 \times 10^{-6}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(K^-\mu^+)/\Gamma_{\text{total}}$		$\Gamma_{14}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;3 \times 10^{-6}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(K_S^0\nu)/\Gamma_{\text{total}}$		$\Gamma_{15}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;2 \times 10^{-5}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

$\Gamma(\bar{p}\pi^+)/\Gamma_{\text{total}}$		$\Gamma_{16}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b><math>&lt;9 \times 10^{-7}</math></b>	90	<sup>1</sup> MCCracken	15	CLAS	$\gamma p \rightarrow K^+\Lambda$

<sup>1</sup> Uses  $B(\Lambda \rightarrow p\pi^-) = (63.9 \pm 0.5)\%$  for normalization mode.

## $\Lambda$ DECAY PARAMETERS

See the “Note on Baryon Decay Parameters” in the neutron Listings. Some early results have been omitted.

### $\alpha_-$ FOR $\Lambda \rightarrow p\pi^-$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.642±0.013 OUR AVERAGE				
0.584±0.046	8500	ASTBURY	75	SPEC
0.649±0.023	10325	CLELAND	72	OSPK
0.67 ± 0.06	3520	DAUBER	69	HBC From Ξ decay
0.645±0.017	10130	OVERSETH	67	OSPK Λ from π <sup>-</sup> ρ
0.62 ± 0.07	1156	CRONIN	63	CNTR Λ from π <sup>-</sup> ρ

### $\alpha_+$ FOR $\bar{\Lambda} \rightarrow \bar{p}\pi^+$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.71 \pm 0.08</math> OUR AVERAGE</b>				
$-0.755 \pm 0.083 \pm 0.063$	$\approx 8.7k$	ABLIKIM	10	BES $J/\psi \rightarrow \Lambda \bar{\Lambda}$
$-0.63 \pm 0.13$	770	TIXIER	88	DM2 $J/\psi \rightarrow \Lambda \bar{\Lambda}$

### $\phi$ ANGLE FOR $\Lambda \rightarrow p\pi^-$ ( $\tan\phi = \beta / \gamma$ )

VALUE (°)	EVTS	DOCUMENT ID	TECN	COMMENT
- 6.5± 3.5 OUR AVERAGE				
- 7.0± 4.5	10325	CLELAND	72	OSPK $\Lambda$ from $\pi^-p$
- 8.0± 6.0	10130	OVERSETH	67	OSPK $\Lambda$ from $\pi^-p$
13.0±17.0	1156	CRONIN	63	OSPK $\Lambda$ from $\pi^-p$

### $\alpha_0 / \alpha_- = \alpha(\Lambda \rightarrow n\pi^0) / \alpha(\Lambda \rightarrow p\pi^-)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
1.01 ±0.07 OUR AVERAGE				
1.000±0.068	4760	<sup>1</sup> OLSEN	70	OSPK $\pi^+ n \rightarrow \Lambda K^+$
1.10 ±0.27		CORK	60	CNTR

<sup>1</sup> OLSEN 70 compares proton and neutron distributions from  $\Lambda$  decay.

### $(\alpha + \bar{\alpha})/(\alpha - \bar{\alpha})$ in $\Lambda \rightarrow p\pi^-, \bar{\Lambda} \rightarrow \bar{p}\pi^+$

Zero if  $CP$  is conserved;  $\alpha_-$  and  $\alpha_+$  are the asymmetry parameters for  $\Lambda \rightarrow p\pi^-$  and  $\bar{\Lambda} \rightarrow \bar{p}\pi^+$  decay. See also the  $\Xi^-$  for a similar test involving the decay chain  $\Xi^- \rightarrow \Lambda\pi^-, \Lambda \rightarrow p\pi^-$  and the corresponding antiparticle chain.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.006±0.021 OUR AVERAGE				
−0.081±0.055±0.059	≈ 8.7k	ABLIKIM	10	BES J/ψ → ΛΛ̄
+0.013±0.022	96k	BARNES	96	CNTR LEAR p̄p → ΛΛ̄
+0.01 ± 0.10	770	TIXIER	88	DM2 J/ψ → ΛΛ̄
−0.02 ± 0.14	10k	<sup>1</sup> CHAUVAT	85	CNTR pp, p̄p ISR
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.07 ± 0.09	4063	BARNES	87	CNTR See BARNES 96

<sup>1</sup> CHAUVAT 85 actually gives  $\alpha_+(\bar{\Lambda})/\alpha_-(\Lambda) = -1.04 \pm 0.29$ . Assumes polarization is same in  $\bar{p}p \rightarrow \bar{\Lambda}X$  and  $pp \rightarrow \Lambda X$ . Tests of this assumption, based on  $C$ -invariance and fragmentation, are satisfied by the data.

### $g_A / g_V$ FOR $\Lambda \rightarrow pe^-\bar{\nu}_e$

Measurements with fewer than 500 events have been omitted. Where necessary, signs have been changed to agree with our conventions, which are given in the “Note on Baryon Decay Parameters” in the neutron Listings. The measurements all assume that the form factor  $g_2 = 0$ . See also the footnote on DWORKIN 90.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>- 0.718 ± 0.015 OUR AVERAGE</b>				
- 0.719 ± 0.016 ± 0.012	37k	<sup>1</sup> DWORKIN	90	SPEC $e\nu$ angular corr.
- 0.70 ± 0.03	7111	BOURQUIN	83	SPEC $\Xi \rightarrow \Lambda \pi^-$
- 0.734 ± 0.031	10k	<sup>2</sup> WISE	81	SPEC $e\nu$ angular correl.
• • • We do not use the following data for averages, fits, limits, etc. • • •				
- 0.63 ± 0.06	817	ALTHOFF	73	OSPK Polarized $\Lambda$

<sup>1</sup> The tabulated result assumes the weak-magnetism coupling  $w \equiv g_W(0)/g_V(0)$  to be 0.97, as given by the CVC hypothesis and as assumed by the other listed measurements. However, DWORKIN 90 *measures*  $w$  to be  $0.15 \pm 0.30$ , and then  $g_A/g_V = -0.731 \pm 0.016$ .

<sup>2</sup> This experiment measures only the absolute value of  $g_A/g_V$ .

## $\Lambda$ REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

MCCRACKEN	15	PR D92 072002	M.E. McCracken <i>et al.</i>	(JLab CLAS Collab.)
ABLIKIM	10	PR D81 012003	M. Ablikim <i>et al.</i>	(BES Collab.)
BARNES	96	PR C54 1877	P.D. Barnes <i>et al.</i>	(CERN PS-185 Collab.)
RYBICKI	96	APP B27 2155	K. Rybicki	
HARTOUNI	94	PRL 72 1322	E.P. Hartouni <i>et al.</i>	(BNL E766 Collab.)
Also		PRL 72 2821 (erratum)	E.P. Hartouni <i>et al.</i>	(BNL E766 Collab.)
LARSON	93	PR D47 799	K.D. Larson <i>et al.</i>	(BNL-811 Collab.)
NOBLE	92	PRL 69 414	A.J. Noble <i>et al.</i>	(BIRM, BOST, BRCO+)
DWORKIN	90	PR D41 780	J. Dworkin <i>et al.</i>	(MICH, WISC, RUTG+)
TIXIER	88	PL B212 523	M.H. Tixier <i>et al.</i>	(DM2 Collab.)
BARNES	87	PL B199 147	P.D. Barnes <i>et al.</i>	(CMU, SACL, LANL+)
BIAGI	86	ZPHY C30 201	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
CHAUVAT	85	PL 163B 273	P. Chauvat <i>et al.</i>	(CERN, CLER, UCLA+)
BOURQUIN	83	ZPHY C21 1	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
COX	81	PRL 46 877	P.T. Cox <i>et al.</i>	(MICH, WISC, RUTG, MINN+)
PONDROM	81	PR D23 814	L. Pondrom <i>et al.</i>	(WISC, MICH, RUTG+)
WISE	81	PL 98B 123	J.E. Wise <i>et al.</i>	(MASA, BNL)
WISE	80	PL 91B 165	J.E. Wise <i>et al.</i>	(MASA, BNL)
SCHACHIN...	78	PRL 41 1348	L. Schachinger <i>et al.</i>	(MICH, RUTG, WISC)
HELLER	77	PL 68B 480	K. Heller <i>et al.</i>	(MICH, WISC, HEIDH)
LINDQUIST	77	PR D16 2104	J. Lindquist <i>et al.</i>	(EFI, OSU, ANL)
Also		JP G2 L211	J. Lindquist <i>et al.</i>	(EFI, WUSL, OSU+)
ZECH	77	NP B124 413	G. Zech <i>et al.</i>	(SIEG, CERN, DORT, HEIDH)
BUNCE	76	PRL 36 1113	G.R.M. Bunce <i>et al.</i>	(WISC, MICH, RUTG)
ASTBURY	75	NP B99 30	P. Astbury <i>et al.</i>	(LOIC, CERN, ETH+)
CLAYTON	75	NP B95 130	E.F. Clayton <i>et al.</i>	(LOIC, RHEL)
ALTHOFF	73	PL 43B 237	K.H. Althoff <i>et al.</i>	(CERN, HEID)
ALTHOFF	73B	NP B66 29	K.H. Althoff <i>et al.</i>	(CERN, HEID)
KATZ	73	Thesis MDDP-TR-74-044	C.N. Katz	(UMD)
POULLARD	73	PL 46B 135	G. Poullard, A. Givernaud, A.C. Borg	(SACL)
BAGGETT	72B	ZPHY 252 362	M.J. Baggett <i>et al.</i>	(HEID)
BAGGETT	72C	PL 42B 379	M.J. Baggett <i>et al.</i>	(HEID)
CLELAND	72	NP B40 221	W.E. Cleland <i>et al.</i>	(CERN, GEVA, LUND)
HYMAN	72	PR D5 1063	L.G. Hyman <i>et al.</i>	(ANL, CMU)
ALTHOFF	71	PL 37B 531	K.H. Althoff <i>et al.</i>	(CERN, HEID)
BALTAY	71B	PR D4 670	C. Baltay <i>et al.</i>	(COLU, BING)
BARONI	71	LNC 2 1256	G. Baroni, S. Pettera, G. Romano	(ROMA)
CANTER	71	PRL 26 868	J. Canter <i>et al.</i>	(STON, COLU)
CANTER	71B	PRL 27 59	J. Canter <i>et al.</i>	(STON, COLU)
DAHL-JENSEN	71	NC 3A 1	E. Dahl-Jensen <i>et al.</i>	(CERN, ANKA, LAUS+)
LINDQUIST	71	PRL 27 612	J. Lindquist <i>et al.</i>	(EFI, WUSL, OSU+)
OLSEN	70	PRL 24 843	S.L. Olsen <i>et al.</i>	(WISC, MICH)
DAUBER	69	PR 179 1252	P.M. Dauber <i>et al.</i>	(LRL)
DOYLE	69	Thesis UCRL 18139	J.C. Doyle	(LRL)
MALONEY	69	PRL 23 425	J.E. Maloney, B. Sechi-Zorn	(UMD)
GRIMM	68	NC 54A 187	H.J. Grimm	(HEID)



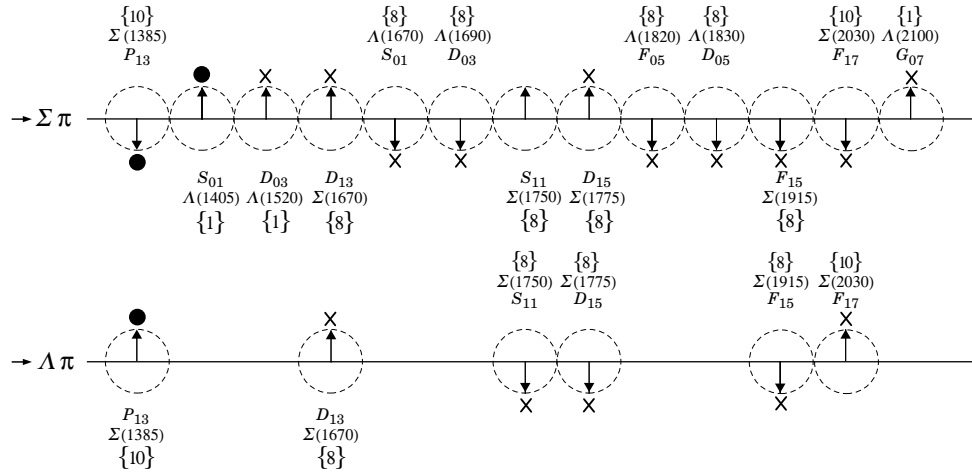
HEPP	68	ZPHY 214 71	V. Hepp, H. Schleich	(HEID)
BADIER	67	PL 25B 152	J. Badier <i>et al.</i>	(EPOL)
MAYEUR	67	U.Libr.Brux.Bul. 32	C. Mayeur, E. Tompa, J.H. Wickens	(BELG, LOUC)
OVERSETH	67	PRL 19 391	O.E. Overseth, R.F. Roth	(MICH, PRIN)
PDG	67	RMP 39 1	A.H. Rosenfeld <i>et al.</i>	(LRL, CERN, YALE)
BURAN	66	PL 20 318	T. Buran <i>et al.</i>	(OSLO)
CHIEN	66	PR 152 1171	C.Y. Chien <i>et al.</i>	(YALE, BNL)
ENGELMANN	66	NC 45A 1038	R. Engelmann <i>et al.</i>	(HEID, REHO)
GIBSON	66	NC 45A 882	W.M. Gibson, K. Green	(BRIS)
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA)
SCHMIDT	65	PR 140B 1328	P. Schmidt	(COLU)
BAGLIN	64	NC 35 977	C. Baglin <i>et al.</i>	(EPOL, CERN, LOUC, RHEL+)
HUBBARD	64	PR 135 B183	J.R. Hubbard <i>et al.</i>	(LRL)
LIND	64	PR 135 B1483	V.G. Lind <i>et al.</i>	(WIS)
RONNE	64	PL 11 357	B.E. Ronne <i>et al.</i>	(CERN, EPOL, LOUC+)
SCHWARTZ	64	Thesis UCRL 11360	J.A. Schwartz	(LRL)
BHOWMIK	63	NC 28 1494	B. Bhowmik, D.P. Goyal	(DELH)
BLOCK	63	PR 130 766	M.M. Block <i>et al.</i>	(NWES, BGNA, SYRA+)
BROWN	63	PR 130 769	J.L. Brown <i>et al.</i>	(LRL, MICH)
CHRETIEN	63	PR 131 2208	M. Chretien <i>et al.</i>	(BRAN, BROW, HARV+)
CRONIN	63	PR 129 1795	J.W. Cronin, O.E. Overseth	(PRIN)
ELY	63	PR 131 868	R.P. Ely <i>et al.</i>	(LRL)
HUMPHREY	62	PR 127 1305	W.E. Humphrey, R.R. Ross	(LRL)
CORK	60	PR 120 1000	B. Cork <i>et al.</i>	(LRL, PRIN, BNL)
CRAWFORD	59B	PRL 2 266	F.S. Crawford <i>et al.</i>	(LRL)

## $\Lambda$ AND $\Sigma$ RESONANCES

**Introduction:** Since our last edition, there have been a few measurements of properties of the lowest  $\Lambda$  and  $\Sigma$  resonances—mostly of masses and widths. But the field remains at a standstill. What follows is a much abbreviated version of the note on  $\Lambda$  and  $\Sigma$  Resonances from our 1990 edition [1]. In particular, see that edition for some representative Argand plots from partial-wave analyses.

the sign depends on conventions used in conjunction with the Clebsch-Gordan coefficients (such as, is the baryon or the meson the “first” particle). If this reaction is partial-wave analyzed and if the overall phase is chosen so that, say, the  $\Sigma(1775)D_{15}$  amplitude at resonance points along the positive imaginary axis (points “up”), then any  $\Sigma$  at resonance will point “up” and any  $\Lambda$  at resonance will point “down” (along the negative imaginary axis). Thus the phase at resonance determines the isospin. The above ignores background amplitudes in the resonating partial waves.

That is the basic idea. In a similar but somewhat more complicated way, the phases of the  $\bar{K}N \rightarrow \Lambda\pi$  and  $\bar{K}N \rightarrow \Sigma\pi$  amplitudes for a resonating wave help determine the SU(3) multiplet to which the resonance belongs. Again, a convention has to be adopted for some overall arbitrary phases: which way is “up”? Our convention is that of Levi-Setti [2] and is shown in Fig. 1, which also compares experimental results with theoretical predictions for the signs of several resonances. In the Listings, a + or – sign in front of a measurement of an inelastic resonance coupling indicates the sign (the *absence* of a sign



**Figure 1.** The signs of the imaginary parts of resonating amplitudes in the  $\bar{K}N \rightarrow \Lambda\pi$  and  $\Sigma\pi$  channels. The signs of the  $\Sigma(1385)$  and  $\Lambda(1405)$ , marked with a ●, are set by convention, and then the others are determined relative to them. The signs required by the SU(3) assignments of the resonances are shown with an arrow, and the experimentally determined signs are shown with an ×.

Table 1 is an attempt to evaluate the status, both overall and channel by channel, of each  $\Lambda$  and  $\Sigma$  resonance in the Particle Listings. The evaluations are of course partly subjective. A blank indicates there is no evidence at all: either the relevant couplings are small or the resonance does not really exist. The main Baryon Summary Table includes only the established resonances (overall status 3 or 4 stars). A number of the 1- and 2-star entries may eventually disappear, but there are certainly many resonances yet to be discovered underlying the established ones.

**Sign conventions for resonance couplings:** In terms of the isospin-0 and -1 elastic scattering amplitudes  $A_0$  and  $A_1$ , the amplitude for  $K^-p \rightarrow \bar{K}^0 n$  scattering is  $\pm(A_1 - A_0)/2$ , where

means that the sign is not determined, *not* that it is positive). For more details, see Appendix II of our 1982 edition [3].

**Errors on masses and widths:** The errors quoted on resonance parameters from partial-wave analyses are often only statistical, and the parameters can change by more than these errors when a different parametrization of the waves is used. Furthermore, the different analyses use more or less the same data, so it is not really appropriate to treat the different determinations of the resonance parameters as independent or to average them together. In any case, the spread of the masses, widths, and branching fractions from the different analyses is certainly a better indication of the uncertainties than are the quoted errors. In the Baryon Summary Table, we usually give a

# Baryon Particle Listings

## $\Lambda$ 's and $\Sigma$ 's, $\Lambda(1405)$

range reflecting the spread of the values rather than a particular value with error.

For three states, the  $\Lambda(1520)$ , the  $\Lambda(1820)$ , and the  $\Sigma(1775)$ , there is enough information to make an overall fit to the various branching fractions. It is then necessary to use the quoted errors, but the errors obtained from the fit should not be taken seriously.

Table 1. The status of the  $\Lambda$  and  $\Sigma$  resonances. Only those with an overall status of \*\*\* or \*\*\*\* are included in the main Baryon Summary Table.

Particle	$J^P$	Overall status	Status as seen in —			
			$N\bar{K}$	$\Lambda\pi$	$\Sigma\pi$	Other channels
$\Lambda(1116)$	1/2+	****		F		$N\pi$ (weakly)
$\Lambda(1405)$	1/2-	****	****	o	****	
$\Lambda(1520)$	3/2-	****	****	r	****	$\Lambda\pi\pi, \Lambda\gamma$
$\Lambda(1600)$	1/2+	***	***	b	**	
$\Lambda(1670)$	1/2-	****	****	i	****	$\Lambda\eta$
$\Lambda(1690)$	3/2-	****	****	d	****	$\Lambda\pi\pi, \Sigma\pi\pi$
$\Lambda(1800)$	1/2-	***	***	d	**	$N\bar{K}^*, \Sigma(1385)\pi$
$\Lambda(1810)$	1/2+	***	***	e	**	$N\bar{K}^*$
$\Lambda(1820)$	5/2+	****	****	n	****	$\Sigma(1385)\pi$
$\Lambda(1830)$	5/2-	****	***	F	****	$\Sigma(1385)\pi$
$\Lambda(1890)$	3/2+	****	****	o	**	$N\bar{K}^*, \Sigma(1385)\pi$
$\Lambda(2000)$		*		r	*	$\Lambda\omega, N\bar{K}^*$
$\Lambda(2020)$	7/2+	*	*	b	*	
$\Lambda(2100)$	7/2-	****	****	i	***	$\Lambda\omega, N\bar{K}^*$
$\Lambda(2110)$	5/2+	***	**	d	*	$\Lambda\omega, N\bar{K}^*$
$\Lambda(2325)$	3/2-	*	*	d		$\Lambda\omega$
$\Lambda(2350)$		***	***	e	*	
$\Lambda(2585)$		**	**	n		
$\Sigma(1193)$	1/2+	****				$N\pi$ (weakly)
$\Sigma(1385)$	3/2+	****		****	****	
$\Sigma(1480)$		*	*	*	*	
$\Sigma(1560)$		**		**	**	
$\Sigma(1580)$	3/2-	*	*	*		
$\Sigma(1620)$	1/2-	**	**	*	*	
$\Sigma(1660)$	1/2+	***	***	*	**	
$\Sigma(1670)$	3/2-	****	****	****	****	several others
$\Sigma(1690)$		**	*	**	*	$\Lambda\pi\pi$
$\Sigma(1750)$	1/2-	***	***	**	*	$\Sigma\eta$
$\Sigma(1770)$	1/2+	*				
$\Sigma(1775)$	5/2-	****	****	****	***	several others
$\Sigma(1840)$	3/2+	*	*	**	*	
$\Sigma(1880)$	1/2+	**	**	**		$N\bar{K}^*$
$\Sigma(1915)$	5/2+	****	***	****	***	$\Sigma(1385)\pi$
$\Sigma(1940)$	3/2-	***	*	***	**	quasi-2-body
$\Sigma(2000)$	1/2-	*		*		$N\bar{K}^*, \Lambda(1520)\pi$
$\Sigma(2030)$	7/2+	****	****	****	**	several others
$\Sigma(2070)$	5/2+	*	*		*	
$\Sigma(2080)$	3/2+	**		**		
$\Sigma(2100)$	7/2-	*		*	*	
$\Sigma(2250)$		***	***	*	*	
$\Sigma(2455)$		**	*	*		
$\Sigma(2620)$		**	*			
$\Sigma(3000)$		*	*	*		
$\Sigma(3170)$		*				multi-body
****	Existence is certain, and properties are at least fairly well explored.					
***	Existence ranges from very likely to certain, but further confirmation is desirable and/or quantum numbers, branching fractions, etc. are not well determined.					
**	Evidence of existence is only fair.					
*	Evidence of existence is poor.					

**Production experiments:** Partial-wave analyses of course separate partial waves, whereas a peak in a cross section or an invariant mass distribution usually cannot be disentangled from background and analyzed for its quantum numbers; and

more than one resonance may be contributing to the peak. Results from partial-wave analyses and from production experiments are generally kept separate in the Listings, and in the Baryon Summary Table results from production experiments are used only for the low-mass states. The  $\Sigma(1385)$  and  $\Lambda(1405)$  of course lie below the  $\bar{K}N$  threshold and nearly everything about them is learned from production experiments; and production and formation experiments agree quite well in the case of  $\Lambda(1520)$  and results have been combined. There is some disagreement between production and formation experiments in the 1600–1700 MeV region: see the note on the  $\Sigma(1670)$ .

## References

1. Particle Data Group, Phys. Lett. **B239**, VIII.64 (1990).
2. R. Levi-Setti, in *Proceedings of the Lund International Conference on Elementary Particles* (Lund, 1969), p. 339.
3. Particle Data Group, Phys. Lett. **111B** (1982).

## $\Lambda(1405) \ 1/2^-$

$I(J^P) = 0(\frac{1}{2}^-)$  Status: \*\*\*\*

The nature of the  $\Lambda(1405)$  has been a puzzle for decades: three-quark state or hybrid; two poles or one. We cannot here survey the rather extensive literature. See, for example, CIEPLY 10, KISSLINGER 11, SEKIYARA 11, and SHEVCHENKO 12A for discussions and earlier references.

It seems to be the universal opinion of the chiral-unitary community that there are two poles in the 1400-MeV region. ZYCHOR 08 presents experimental evidence against the two-pole model, but this is disputed by GENG 07A. See also REVAI 09, which finds little basis for choosing between one- and two-pole models; and IKEDA 12, which favors the two-pole model.

A single, ordinary three-quark  $\Lambda(1405)$  fits nicely into a  $J^P = 1/2^-$   $SU(4)$   $\bar{4}$  multiplet, whose other members are the  $\Lambda_c(2595)^+$ ,  $\Xi_c(2790)^+$ , and  $\Xi_c(2790)^0$ ; see Fig. 1 of our note on “Charmed Baryons.”

## POLE STRUCTURE OF THE $\Lambda(1405)$ REGION

Written November 2015 by Ulf-G. Meißner (Bonn Univ. / FZ Jülich) and Tetsuo Hyodo (YITP, Kyoto Univ.).

The  $\Lambda(1405)$  resonance emerges in the meson-baryon scattering amplitude with the strangeness  $S = -1$  and isospin  $I = 0$ . It is the archetype of what is called a dynamically generated resonance, as pioneered by Dalitz and Tuan [1]. The most powerful and systematic approach for the low-energy regime of the strong interactions is chiral perturbation theory (ChPT), see e.g. Ref. 2. A perturbative calculation is, however, not applicable to this sector because of the existence of the  $\Lambda(1405)$  just below the  $\bar{K}N$  threshold. In this case, ChPT has to be combined with a non-perturbative resummation technique, just as in the case of the nuclear forces. By solving the Lippmann-Schwinger equation with the interaction kernel determined by ChPT and using a particular regularization, in Ref. 3 a successful description of the low-energy  $K^-p$  scattering data as well as the mass distribution of the  $\Lambda(1405)$  was achieved (for further developments, see Ref. 4 and references therein).

The study of the pole structure was initiated by Ref. 5, which finds two poles of the scattering amplitude in the complex energy plane between the  $\bar{K}N$  and  $\pi\Sigma$  thresholds. The spectrum in experiments exhibits one effective resonance shape,

while the existence of two poles results in the reaction-dependent lineshape [6]. The origin of this two-pole structure is attributed to the two attractive channels of the leading order interaction in the SU(3) basis (singlet and octet) [6] and in the isospin basis ( $\bar{K}N$  and  $\pi\Sigma$ ) [7]. It is remarkable that the sign and the strength of the leading order interaction is determined by a low-energy theorem of chiral symmetry, i.e. the so-called Weinberg-Tomozawa term. The two-pole nature of the  $\Lambda(1405)$  is qualitatively different from the case of the N(1440) resonance. Two poles of the N(1440) appear on different Riemann sheets of the complex energy plane separated by the  $\pi\Delta$  branch point. These poles reflect a single state, with a nearby pole and a more distant shadow pole. In contrast, the two poles in the  $\Lambda(1405)$  region on the same Riemann sheet (where  $\pi\Sigma$  channels are unphysical and all other channels physical, correspondingly to the one, connected to the real axis between the  $\pi\Sigma$  and  $\bar{K}N$  thresholds) are generated from two attractive forces mentioned above [6,7].

Recently, various new experimental results on the  $\Lambda(1405)$  have become available [4]. Among these, the most striking measurement is the precise determination of the energy shift and width of kaonic hydrogen by the SIDDHARTA collaboration [8], [9], which provides a quantitative and stringent constraint on the  $K^-p$  amplitude at threshold through the improved Deser formula [10]. Systematic studies with error analyses based on the next-to-leading order ChPT interaction including the SIDDHARTA constraint have been performed by various groups [11–15]. All these studies confirm that the new kaonic hydrogen data are compatible with the scattering data above threshold.

**Table 1:** Comparison of the pole positions of  $\Lambda(1405)$  in the complex energy plane from next-to-leading order chiral unitary coupled-channel approaches including the SIDDHARTA constraint.

approach	pole 1 [MeV]	pole 2 [MeV]
Refs. 11,12, NLO	$1424^{+7}_{-23} - i 26^{+3}_{-14}$	$1381^{+18}_{-6} - i 81^{+19}_{-8}$
Ref. 14, Fit II	$1421^{+3}_{-2} - i 19^{+8}_{-5}$	$1388^{+9}_{-9} - i 114^{+24}_{-25}$
Ref. 15, solution #2	$1434^{+2}_{-2} - i 10^{+2}_{-1}$	$1330^{+4}_{-5} - i 56^{+17}_{-11}$
Ref. 15, solution #4	$1429^{+8}_{-7} - i 12^{+2}_{-3}$	$1325^{+15}_{-15} - i 90^{+12}_{-18}$

The results of the pole positions of  $\Lambda(1405)$  in the various approaches are summarized in Table 1. We may regard the difference among the calculations as a systematic error, which stems from the various approximations of the Bethe-Salpeter equation, the fitting procedure, and also the inclusion of SU(3) breaking effects such as the choice of the various meson decay constants, and so on. The main component for the  $\Lambda(1405)$  is the pole 1, whose position converges within a relatively small region near the  $\bar{K}N$  threshold. On the other hand, the position of the pole 2 shows a sizeable scatter. Detailed studies of the  $\pi\Sigma$  spectrum in various reaction processes, together with

the precise experimental lineshape (see e.g. the recent precise photoproduction data from the LEPS collaboration [16] and from the CLAS collaboration [17,18], electroproduction data from the CLAS collaboration [19], and proton-proton collision data from COSY [20] and the HADES collaboration [21]), will shed light on the position of the second pole. The  $\pi\Sigma$  spectra from the CLAS data and the HADES data are analyzed in Ref. 22 and Ref. 23, respectively. Although the result of the pole positions in Ref. 22 is similar to those in Table 1, the pole found in Ref. 23 is not compatible with other results. Therefore, the analysis with only the  $\pi\Sigma$  spectrum is not completely conclusive. It is thus desirable to perform a comprehensive analysis of  $\pi\Sigma$  spectra together with the systematic error analysis of the scattering data as done in Ref. 15. It was shown there that several solutions, which agree with the scattering data are ruled out, if confronted with the recent CLAS data. The remaining solutions are collected as solution #2 and solution #4 of Ref. 15 in Table 1.

## References

1. R.H. Dalitz, S.F. Tuan Phys. Rev. Lett. **2**, 425 (1959).
2. V. Bernard *et al.*, Int. J. Mod. Phys. **E4**, 193 (1995).
3. N. Kaiser *et al.*, Nucl. Phys. **A594**, 325 (1995).
4. T. Hyodo, D. Jido, Prog. in Part. Nucl. Phys. **67**, 55 (2012).
5. J.A. Oller, U.-G. Meißner, Phys. Lett. **B500**, 263 (2001).
6. D. Jido *et al.*, Nucl. Phys. **A725**, 181 (2003).
7. T. Hyodo, W. Weise, Phys. Rev. **C77**, 035204 (2008).
8. M. Bazzi *et al.*, Phys. Lett. **B704**, 113 (2011).
9. M. Bazzi *et al.*, Nucl. Phys. **A881**, 88 (2012).
10. U.-G. Meißner *et al.*, Eur. Phys. J. **C35**, 349 (2004).
11. Y. Ikeda *et al.*, Phys. Lett. **B706**, 63 (2011).
12. Y. Ikeda *et al.*, Nucl. Phys. **A881**, 98 (2012).
13. M. Mai, U.-G. Meißner, Nucl. Phys. **A900**, 51 (2013).
14. Z.-H. Guo, J. Oller, Phys. Rev. **C87**, 035202 (2013).
15. M. Mai, U.-G. Meißner, Eur. Phys. J. **A51**, 30 (2015).
16. M. Niiyama *et al.*, Phys. Rev. **C78**, 035202 (2008).
17. K. Moriya *et al.*, Phys. Rev. **C87**, 035206 (2013).
18. K. Moriya *et al.*, Phys. Rev. Lett. **112**, 082004 (2014).
19. H.Y. Lu *et al.*, Phys. Rev. **C88**, 045202 (2013).
20. I. Zychor *et al.*, Phys. Lett. **B660**, 167 (2008).
21. G. Agakishiev *et al.*, Phys. Rev. **C87**, 025201 (2013).
22. L. Roca, E. Oset, Phys. Rev. **C87**, 055201 (2013).
23. M. Hassanvand *et al.*, Phys. Rev. **C87**, 055202 (2013).

## $\Lambda(1405)$ MASS

### PRODUCTION EXPERIMENTS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1405.1^{+1.3}_{-1.0}</math> OUR AVERAGE</b>				
$1405^{+11}_{-9}$		HASSANVAND 13	SPEC	$pp \rightarrow p\Lambda(1405)K^+$
$1405^{+1.4}_{-1.0}$		ESMAILI	10	RVUE $^4\text{He } K^- \rightarrow \Sigma^\pm \pi^\mp X$ at rest
$1406.5 \pm 4.0$		<sup>1</sup> DALITZ	91	M-matrix fit

# Baryon Particle Listings

## $\Lambda(1405)$ , $\Lambda(1520)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1391 $\pm$ 1	700	<sup>1</sup> HEMINGWAY	85	HBC	$K^- p$ 4.2 GeV/ <i>c</i>
$\sim$ 1405	400	<sup>2</sup> THOMAS	73	HBC	$\pi^- p$ 1.69 GeV/ <i>c</i>
1405	120	BARBARO...	68B	DBC	$K^- d$ 2.1–2.7 GeV/ <i>c</i>
1400 $\pm$ 5	67	BIRMINGHAM	66	HBC	$K^- p$ 3.5 GeV/ <i>c</i>
1382 $\pm$ 8		ENGLER	65	HDBC	$\pi^- p, \pi^+ d$ 1.68 GeV/ <i>c</i>
1400 $\pm$ 24		MUSGRAVE	65	HBC	$\bar{p} p$ 3–4 GeV/ <i>c</i>
1410		ALEXANDER	62	HBC	$\pi^- p$ 2.1 GeV/ <i>c</i>
1405		ALSTON	62	HBC	$K^- p$ 1.2–0.5 GeV/ <i>c</i>
1405		ALSTON	61B	HBC	$K^- p$ 1.15 GeV/ <i>c</i>

### EXTRAPOLATIONS BELOW $N\bar{K}$ THRESHOLD

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1407.56 or 1407.50		<sup>3</sup> KIMURA	00	potential model
1411		<sup>4</sup> MARTIN	81	K-matrix fit
1406		<sup>5</sup> CHAO	73	DPWA 0-range fit (sol. B)
1421		MARTIN	70	RVUE Constant K-matrix
1416 $\pm$ 4		MARTIN	69	HBC Constant K-matrix
1403 $\pm$ 3		KIM	67	HBC K-matrix fit
1407.5 $\pm$ 1.2		<sup>6</sup> KITTEL	66	0-effective-range fit
1410.7 $\pm$ 1.0		KIM	65	HBC 0-effective-range fit
1409.6 $\pm$ 1.7		<sup>6</sup> SAKITT	65	HBC 0-effective-range fit

### $\Lambda(1405)$ WIDTH

### PRODUCTION EXPERIMENTS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>50.5<math>\pm</math> 2.0 OUR AVERAGE</b>				
62 $\pm$ 10		HASSANVAND 13	SPEC	$pp \rightarrow p\Lambda(1405) K^+$
50 $\pm$ 2		<sup>1</sup> DALITZ	91	M-matrix fit
• • • We do not use the following data for averages, fits, limits, etc. • • •				
24 $+ \begin{smallmatrix} 4 \\ - 3 \end{smallmatrix}$		ESMAILI	10	RVUE $^4\text{He } K^- \rightarrow \Sigma^\pm \pi^\mp X$ at rest
32 $\pm$ 1	700	<sup>1</sup> HEMINGWAY	85	HBC $K^- p$ 4.2 GeV/ <i>c</i>
45 to 55	400	<sup>2</sup> THOMAS	73	HBC $\pi^- p$ 1.69 GeV/ <i>c</i>
35	120	BARBARO...	68B	DBC $K^- d$ 2.1–2.7 GeV/ <i>c</i>
50 $\pm$ 10	67	BIRMINGHAM	66	HBC $K^- p$ 3.5 GeV/ <i>c</i>
89 $\pm$ 20		ENGLER	65	HDBC
60 $\pm$ 20		MUSGRAVE	65	HBC
35 $\pm$ 5		ALEXANDER	62	HBC
50		ALSTON	62	HBC
20		ALSTON	61B	HBC

### EXTRAPOLATIONS BELOW $N\bar{K}$ THRESHOLD

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
50.24 or 50.26		<sup>3</sup> KIMURA	00	potential model
30		<sup>4</sup> MARTIN	81	K-matrix fit
55		<sup>5,7</sup> CHAO	73	DPWA 0-range fit (sol. B)
20		MARTIN	70	RVUE Constant K-matrix
29 $\pm$ 6		MARTIN	69	HBC Constant K-matrix
50 $\pm$ 5		KIM	67	HBC K-matrix fit
34.1 $\pm$ 4.1		<sup>6</sup> KITTEL	66	HBC
37.0 $\pm$ 3.2		KIM	65	HBC
28.2 $\pm$ 4.1		<sup>6</sup> SAKITT	65	HBC

### $\Lambda(1405)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Sigma \pi$	100 %
$\Gamma_2 \quad \Lambda \gamma$	
$\Gamma_3 \quad \Sigma^0 \gamma$	
$\Gamma_4 \quad N \bar{K}$	

### $\Lambda(1405)$ PARTIAL WIDTHS

#### $\Gamma(\Lambda \gamma)$

VALUE (keV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
27 $\pm$ 8	BURKHARDT 91	Isobar model fit

#### $\Gamma(\Sigma^0 \gamma)$

VALUE (keV)	DOCUMENT ID	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •		
10 $\pm$ 4 or 23 $\pm$ 7	BURKHARDT 91	Isobar model fit

### $\Lambda(1405)$ BRANCHING RATIOS

#### $\Gamma(N\bar{K})/\Gamma(\Sigma \pi)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3	95	HEMINGWAY 85	HBC	$K^- p$ 4.2 GeV/ <i>c</i>

### $\Lambda(1405)$ FOOTNOTES

- <sup>1</sup> DALITZ 91 fits the HEMINGWAY 85 data.
- <sup>2</sup> THOMAS 73 data is fit by CHAO 73 (see next section).
- <sup>3</sup> The KIMURA 00 values are from fits A and B from a coupled-channel potential model using low-energy  $\bar{K} N$  and  $\Sigma \pi$  data, kaonic-hydrogen x-ray measurements, and our  $\Lambda(1405)$  mass and width. The results bear mainly on the *nature* of the  $\Lambda(1405)$ : three-quark state or  $\bar{K} N$  bound state.
- <sup>4</sup> The MARTIN 81 fit includes the  $K^\pm p$  forward scattering amplitudes and the dispersion relations they must satisfy.
- <sup>5</sup> See also the accompanying paper of THOMAS 73.
- <sup>6</sup> Data of SAKITT 65 are used in the fit by KITTEL 66.
- <sup>7</sup> An asymmetric shape, with  $\Gamma/2 = 41$  MeV below resonance, 14 MeV above.

### $\Lambda(1405)$ REFERENCES

HASSANVAND 13	PR C87 055202	M. Hassanvand <i>et al.</i>
Also	PR C88 019905 (errat.)	M. Hassanvand <i>et al.</i>
IKEDA 12	NP A881 98	Y. Ikeda, T. Hyodo, W. Weise (TINT, RIKEN, MUNT)
SHEVCHENKO 12A	PR C85 034001	N.V. Shevchenko (REZ)
KISSLINGER 11	EPJ A47 8	L.S. Kisslinger, E.M. Henley (CMU, WASH)
SEKIHARA 11	PR C83 055202	T. Sekihara, T. Hyodo, D. Jido (KYOT, KYOTU+)
CIERLY 10	EPJ A43 191	A. Cierly, J. Smekal (NPI, Tech. U. Czech Rep.)
ESMAILI 10	PL B686 23	J. Esmaili, Y. Akaishi, T. Yamazaki (RIKEN, ISUT+)
REVAI 09	PR C79 035202	J. Revai, N.V. Shevchenko (BUDA, NPI Czech Rep.)
ZYCHOR 08	PL B660 167	I. Zychor <i>et al.</i> (COSY-ANKE Collab.)
GENG 07A	EPJ A34 405	L.S. Geng, E. Oset (VALE)
KIMURA 00	PR C62 015206	M. Kimura <i>et al.</i>
BURKHARDT 91	PR C44 607	H. Burkhardt, J. Lowe (NOTT, UNM, BIRM)
DALITZ 91	JP G17 289	R.H. Dalitz, A. Deloff (OXFTP, WINR)
HEMINGWAY 85	NP B253 742	R.J. Hemingway (CERN, J)
MARTIN 81	NP B179 33	A.D. Martin (DURH)
CHAO 73	NP B56 46	Y.A. Chao <i>et al.</i> (RHEL, CMU, LOUC)
THOMAS 73	NP B56 15	D.W. Thomas <i>et al.</i> (CMU, J)
MARTIN 70	NP B16 479	A.D. Martin, G.G. Ross (DURH)
MARTIN 69	PR 183 1352	B.R. Martin, M. Sakitt (LOUC, BNL)
Also	PR 183 1345	B.R. Martin, M. Sakitt (LOUC, BNL)
BARBARO... 68B	PRL 21 573	A. Barbaro-Galtieri <i>et al.</i> (LRL, SLAC)
KIM 67	PRL 19 1074	J.K. Kim (YALE)
BIRMINGHAM 66	PR 152 1148	M. Haque <i>et al.</i> (BIRM, GLAS, LOIC, OXF+)
KITTEL 66	PL 21 349	W. Kittel, G. Otter, I. Wacek (VIEN)
ENGLER 65	PRL 15 224	A. Engler <i>et al.</i> (CMU, BNL, IJ)
KIM 65	PRL 14 29	J.K. Kim (COLU)
MUSGRAVE 65	NC 35 735	B. Musgrave <i>et al.</i> (BIRM, CERN, EPOL+)
SAKITT 65	PR 139 8719	M. Sakitt <i>et al.</i> (UMD, LRL)
ALEXANDER 62	PRL 8 447	G. Alexander <i>et al.</i> (LRL)
ALSTON 62	CERN Conf. 311	M.H. Alston <i>et al.</i> (LRL)
ALSTON 61B	PRL 6 698	M.H. Alston <i>et al.</i> (LRL)

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BATTY 89	NC 102A 255	C.J. Batty, A. Gal (RAL, HEBR)
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SCHNICK 87	PRL 58 1719	J. Schinick, R.H. Landau (ORST)
CAPSTICK 86	PR D34 2809	S. Capstick, N. Isgur (TNTO)
JENNINGS 86	PL B176 229	B.K. Jennings (TRIU)
MALTMAN 86	PR D34 1372	K. Maltman, N. Isgur (LANL, TNTO)
ZHONG 86	PL B171 471	Y.S. Zhong <i>et al.</i> (ADLD, TRIU, SURR)
BURKHARDT 85	NP A440 653	H. Burkhardt, J. Lowe, A.S. Rosenthal (NOTT+)
DAREWYCH 85	PR D32 1765	J.W. Darewych, R. Koniuk, N. Isgur (YORKC, TNTO)
VEIT 85	PR D31 1033	E.A. Veit <i>et al.</i> (TRIU, ADLD, SURR)
KIANG 84	PR C30 1638	D. Kiang <i>et al.</i> (DALH, MCMS)
MILLER 84	Conference paper	D.J. Miller (LOUC)
Conf. Intersections between Particle and Nuclear Physics, p. 783		
VANDIJK 84	PR D30 937	W. van Dijk (MCMS)
VEIT 84	PL 137B 415	E.A. Veit <i>et al.</i> (TRIU, SURR, CERN)
DALITZ 82	Heid. Conf.	R.H. Dalitz <i>et al.</i> (OXFTP)
Heidelberg Conf., p. 201		
DALITZ 81	Kaon Conf.	R.H. Dalitz, J.G. McGinley (OXFTP)
Low and Intermediate Energy Kaon-Nucleon Physics, p.381		
MARTIN 81B	Kaon Conf.	A.D. Martin (DURH)
Low and Intermediate Energy Kaon-Nucleon Physics, p. 97		
OADES 77	NC 42A 462	G.C. Oades, G. Rasche (AARH, ZURI)
SHAW 72	Purdue Conf. 417	G.L. Shaw (UCI)
BARBARO... 72	LBL-555	A. Barbaro-Galtieri (LBL)
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RAJASEKA... 72	PR D5 610	G. Rajasekaran (TATA)
Earlier papers also cited in RAJASEKARAN 72.		
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MARTIN 71	PL 35B 62	A.D. Martin, A.D. Martin, G.G. Ross (DURH, LOUC+)
DALITZ 67	PR 153 1617	R.H. Dalitz, T.C. Wong, G. Rajasekaran (OXFTP+)
DONALD 66	PL 22 711	R.A. Donald <i>et al.</i> (LIVP)
KADYK 66	PRL 17 599	J.A. Kadyk <i>et al.</i> (LRL)
ABRAMS 65	PR 139 B454	G.S. Abrams, B. Sechi-Zorn (UMD)

## $\Lambda(1520) \ 3/2^-$

$$I(J^P) = 0(\frac{3}{2}^-) \text{ Status: } ***$$

Discovered by FERRO-LUZZI 62; the elaboration in WATSON 63 is the classic paper on the Breit-Wigner analysis of a multichannel resonance.

The measurements of the mass, width, and elasticity published before 1975 are now obsolete and have been omitted. They were last listed in our 1982 edition Physics Letters **111B** 1 (1982).

Production and formation experiments agree quite well, so they are listed together here.

See key on page 601

# Baryon Particle Listings

## $\Lambda(1520)$

 **$\Lambda(1520)$  MASS**

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1519.5 ±1.0 OUR ESTIMATE</b>				
<b>1519.54 ±0.17 OUR AVERAGE</b>				
1519.6 ±0.5		ZHANG	13A DPWA	Multichannel
1520.4 ±0.6 ±1.5		QIANG	10 SPEC	$ep \rightarrow e' K^+ X$ (fit to $X$ )
1517.3 ±1.5	300	BARBER	80D SPEC	$\gamma p \rightarrow \Lambda(1520) K^+$
1517.8 ±1.2	5k	BARLAG	79 HBC	$K^- p$ 4.2 GeV/c
1520.0 ±0.5		ALSTON-...	78 DPWA	$\bar{K} N \rightarrow \bar{K} N$
1519.7 ±0.3	4k	CAMERON	77 HBC	$K^- p$ 0.96–1.36 GeV/c
1519 ±1		GOPAL	77 DPWA	$\bar{K} N$ multichannel
1519.4 ±0.3	2000	CORDEN	75 DBC	$K^- d$ 1.4–1.8 GeV/c

 **$\Lambda(1520)$  WIDTH**

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>15.6 ±1.0 OUR ESTIMATE</b>				
<b>15.73 ±0.29 OUR AVERAGE</b>				Error includes scale factor of 1.1.
17 ±1		ZHANG	13A DPWA	Multichannel
18.6 ±1.9 ±1.0		QIANG	10 SPEC	$ep \rightarrow e' K^+ X$ (fit to $X$ )
16.3 ±3.3	300	BARBER	80D SPEC	$\gamma p \rightarrow \Lambda(1520) K^+$
16 ±1		GOPAL	80 DPWA	$\bar{K} N \rightarrow \bar{K} N$
14 ±3	677	BARLAG	79 HBC	$K^- p$ 4.2 GeV/c
15.4 ±0.5		ALSTON-...	78 DPWA	$\bar{K} N \rightarrow \bar{K} N$
16.3 ±0.5	4k	CAMERON	77 HBC	$K^- p$ 0.96–1.36 GeV/c
15.0 ±0.5		GOPAL	77 DPWA	$\bar{K} N$ multichannel
15.5 ±1.6	2000	CORDEN	75 DBC	$K^- d$ 1.4–1.8 GeV/c

<sup>1</sup> From the best-resolution sample of  $\Lambda\pi\pi$  events only. **$\Lambda(1520)$  POLE POSITION****REAL PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
...			...
1518	ZHANG	13A DPWA	Multichannel
1518.8	QIANG	10 SPEC	$ep \rightarrow e' K^+ X$ (fit to $X$ )

**−2×IMAGINARY PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
...			...
16	ZHANG	13A DPWA	Multichannel
17.2	QIANG	10 SPEC	$ep \rightarrow e' K^+ X$ (fit to $X$ )

 **$\Lambda(1520)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	(45 ±1) %
$\Gamma_2$ $\Sigma\pi$	(42 ±1) %
$\Gamma_3$ $\Lambda\pi\pi$	(10 ±1) %
$\Gamma_4$ $\Sigma(1385)\pi$	
$\Gamma_5$ $\Sigma(1385)\pi(\rightarrow \Lambda\pi\pi)$	
$\Gamma_6$ $\Lambda(\pi\pi)s$ -wave	
$\Gamma_7$ $\Sigma\pi\pi$	(0.9 ±0.1) %
$\Gamma_8$ $\Lambda\gamma$	(0.85 ±0.15) %
$\Gamma_9$ $\Sigma^0\gamma$	

**CONSTRAINED FIT INFORMATION**

An overall fit to 9 branching ratios uses 28 measurements and one constraint to determine 6 parameters. The overall fit has a  $\chi^2 = 18.9$  for 23 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−63				
$x_3$	−32	−34			
$x_7$	−4	−3	−1		
$x_8$	−8	−7	−3	0	
$x_9$	−24	−21	−10	−1	−1
	$x_1$	$x_2$	$x_3$	$x_7$	$x_8$

 **$\Lambda(1520)$  BRANCHING RATIOS**

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

 **$\Gamma(N\bar{K})/\Gamma_{\text{total}}$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.45 ±0.01 OUR ESTIMATE</b>			
<b>0.448 ±0.007 OUR FIT</b>			Error includes scale factor of 1.2.
<b>0.456 ±0.010 OUR AVERAGE</b>			
0.47 ±0.04	ZHANG	13A DPWA	Multichannel
0.47 ±0.02	GOPAL	80 DPWA	$\bar{K} N \rightarrow \bar{K} N$
0.45 ±0.03	ALSTON-...	78 DPWA	$\bar{K} N \rightarrow \bar{K} N$
0.448 ±0.014	CORDEN	75 DBC	$K^- d$ 1.4–1.8 GeV/c
...			...
0.47 ±0.01	GOPAL	77 DPWA	See GOPAL 80
0.42	MAST	76 HBC	$K^- p \rightarrow \bar{K}^0 n$

 **$\Gamma_1/\Gamma$**  **$\Gamma(\Sigma\pi)/\Gamma_{\text{total}}$** 

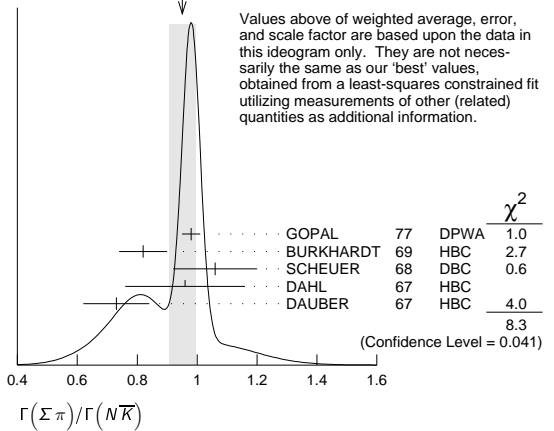
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.42 ±0.01 OUR ESTIMATE</b>			
<b>0.421 ±0.007 OUR FIT</b>			Error includes scale factor of 1.2.
<b>0.425 ±0.011 OUR AVERAGE</b>			
0.47 ±0.05	ZHANG	13A DPWA	Multichannel
0.426 ±0.014	CORDEN	75 DBC	$K^- d$ 1.4–1.8 GeV/c
0.418 ±0.017	BARBARO-...	69B HBC	$K^- p$ 0.28–0.45 GeV/c
...			...
0.46	KIM	71 DPWA	K-matrix analysis

 **$\Gamma_2/\Gamma$**  **$\Gamma(\Sigma\pi)/\Gamma(N\bar{K})$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.940 ±0.026 OUR FIT</b>			Error includes scale factor of 1.3.
<b>0.95 ±0.04 OUR AVERAGE</b>			Error includes scale factor of 1.7. See the ideogram below.
0.98 ±0.03	<sup>2</sup> GOPAL	77 DPWA	$\bar{K} N$ multichannel
0.82 ±0.08	BURKHARDT	69 HBC	$K^- p$ 0.8–1.2 GeV/c
1.06 ±0.14	SCHEUER	68 DBC	$K^- N$ 3 GeV/c
0.96 ±0.20	DAHL	67 HBC	$\pi^- p$ 1.6–4 GeV/c
0.73 ±0.11	DAUBER	67 HBC	$K^- p$ 2 GeV/c
...			...
1.06 ±0.12	BERTHON	74 HBC	Quasi-2-body $\sigma$
1.72 ±0.78	MUSGRAVE	65 HBC	

<sup>2</sup> The  $\bar{K} N \rightarrow \Sigma\pi$  amplitude at resonance is  $+0.46 \pm 0.01$ .

WEIGHTED AVERAGE  
0.95±0.04 (Error scaled by 1.7)

 **$\Gamma(\Lambda\pi\pi)/\Gamma_{\text{total}}$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.10 ±0.01 OUR ESTIMATE</b>			
<b>0.095 ±0.005 OUR FIT</b>			Error includes scale factor of 1.2.
<b>0.096 ±0.008 OUR AVERAGE</b>			Error includes scale factor of 1.6.
0.091 ±0.006	CORDEN	75 DBC	$K^- d$ 1.4–1.8 GeV/c
0.11 ±0.01	<sup>3</sup> MAST	73B IPWA	$K^- p \rightarrow \Lambda\pi\pi$

<sup>3</sup> Assumes  $\Gamma(N\bar{K})/\Gamma_{\text{total}} = 0.46 \pm 0.02$ . **$\Gamma(\Lambda\pi\pi)/\Gamma(N\bar{K})$** 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.212 ±0.012 OUR FIT</b>			Error includes scale factor of 1.2.
<b>0.202 ±0.021 OUR AVERAGE</b>			
0.22 ±0.03	BURKHARDT	69 HBC	$K^- p$ 0.8–1.2 GeV/c
0.19 ±0.04	SCHEUER	68 DBC	$K^- N$ 3 GeV/c
0.17 ±0.05	DAHL	67 HBC	$\pi^- p$ 1.6–4 GeV/c
0.21 ±0.18	DAUBER	67 HBC	$K^- p$ 2 GeV/c
...			...
0.27 ±0.13	BERTHON	74 HBC	Quasi-2-body $\sigma$
0.2	KIM	71 DPWA	K-matrix analysis

# Baryon Particle Listings

## $\Lambda(1520)$ , $\Lambda(1600)$

$\Gamma(\Sigma\pi)/\Gamma(\Lambda\pi\pi)$				$\Gamma_2/\Gamma_3$
VALUE		DOCUMENT ID	TECN	COMMENT
<b><math>4.43\pm0.25</math> OUR FIT</b>	Error includes	scale factor of 1.2.		
<b><math>3.9 \pm 0.6</math> OUR AVERAGE</b>				

3.9 $\pm$ 1.0	UHLIG	67	HBC	$K^-p$ 0.9–1.0 GeV/c
3.3 $\pm$ 1.1	BIRMINGHAM	66	HBC	$K^-p$ 3.5 GeV/c
4.5 $\pm$ 1.0	ARMENTEROS65c		HBC	

$\Gamma(\Sigma(1385)\pi)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.041\pm0.005</math></b>	CHAN	72	HBC	$K^-p \rightarrow \Lambda\pi\pi$

$\Gamma(\Sigma(1385)\pi(\rightarrow\Lambda\pi\pi))/\Gamma(\Lambda\pi\pi)$	$\Gamma_5/\Gamma_3$
The $\Lambda\pi\pi$ mode is largely due to $\Sigma(1385)\pi$ . Only the values of $(\Sigma(1385)\pi)/(\Lambda 2\pi)$ given by MAST 73b and CORDEN 75 are based on real 3-body partial-wave analyses. The discrepancy between the two results is essentially due to the different hypotheses made concerning the shape of the $(\pi\pi)_{\text{S-wave}}$ state.	
VALUE	CL%
0.58 $\pm$ 0.22	
0.82 $\pm$ 0.10	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
<0.44	90
0.39 $\pm$ 0.10	
4 Both $\Sigma(1385)\pi DS_{03}$ and $\Sigma(\pi\pi) DP_{03}$ contribute.	
5 The central bin (1514–1524 MeV) gives $0.74 \pm 0.10$ ; other bins are lower by 2-to-5 standard deviations.	

$\Gamma(\Lambda(\pi\pi)_{\text{S-wave}})/\Gamma(\Lambda\pi\pi)$	$\Gamma_6/\Gamma_3$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.20\pm0.08</math></b>	CORDEN	75	DBC	$K^-d$ 1.4–1.8 GeV/c

<0.44	90	WIELAND	11	SPHR	$\gamma p \rightarrow K^+ \Lambda(1520)$
$0.39 \pm 0.10$	<sup>5</sup>	BURKHARDT	71	HBC	$K^- p \rightarrow (\Lambda \pi \pi) \pi$
<sup>4</sup> Both $\Sigma(1385) \pi D S_03$ and $\Sigma(\pi \pi) D P_03$ contribute.					
<sup>5</sup> The central bin (1514–1524 MeV) gives $0.74 \pm 0.10$ ; other bins are lower by 2-to-5 standard deviations.					

$\Gamma(\Lambda\gamma)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.20\pm0.08</math></b>	CORDEN	75	DBC	$K^-d$ 1.4–1.8 GeV/c

$\Gamma(\Sigma\pi\pi)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.009 \pm 0.001</math> OUR ESTIMATE</b>				
<b><math>0.0086 \pm 0.0005</math> OUR FIT</b>				
<b><math>0.0086 \pm 0.0005</math> OUR AVERAGE</b>				
$0.007 \pm 0.002$	<sup>6</sup> CORDEN	75	DBC	$K^-d$ 1.4–1.8 GeV/c
$0.0085 \pm 0.0006$	<sup>7</sup> MAST	73	MPWA	$K^-p \rightarrow \Sigma\pi\pi$
$0.010 \pm 0.0015$	BARBARO...	69b	HBC	$K^-p$ 0.28–0.45 GeV/c
<sup>6</sup> Much of the $\Sigma\pi\pi$ decay proceeds via $\Sigma(1385)\pi$ .				
<sup>7</sup> Assumes $\Gamma(N\bar{K})/\Gamma_{\text{total}} = 0.46$ .				

$\Gamma(\Lambda\gamma)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma$			
<i>VALUE (units <math>10^{-3}</math>)</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>	
<b>8.5±1.5 OUR ESTIMATE</b>				
<b>8.8±1.1 OUR FIT</b>				
<b>8.8±1.1 OUR AVERAGE</b>				
10.7±2.9 <sup>+1.5</sup> <sub>-0.4</sub>	32	TAYLOR	05	CLAS $\gamma p \rightarrow K^+ \Lambda \gamma$
10.2±2.1±1.5	290	ANTIPOV	04A	SPNX $p N(C) \rightarrow \Lambda(1520) K^+ N(C)$
8.0±1.4	238	MAST	68B	HBC Using $\Gamma(N\overline{K})/\Gamma_{\text{total}} = 0.45$

$\Gamma(\Sigma^0\gamma)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.0193 \pm 0.0034</math> OUR FIT</b>				
<b><math>0.02 \pm 0.0035</math></b>	<sup>8</sup> MAST	68B	HBC	Not measured; see note
<sup>8</sup> Calculated from $\Gamma(\Lambda\gamma)/\Gamma_{\text{total}}$ , assuming SU(3). Needed to constrain the sum of all the branching ratios to be unity.				

### $\Lambda(1520)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
WIELAND	11	EPJ A47 47	F. Wieland <i>et al.</i>	(ELSA SAPHIR Collab.)
QIANG	10	PL B694 123	Y. Qiang <i>et al.</i>	(DUKE, JEFF, PNPI, GWU+)
TAYLOR	05	PR C71 054609	S. Taylor <i>et al.</i>	(JLab CLAS Collab.)
Also		PR C72 039902 (errat.)	S. Taylor <i>et al.</i>	(JLab CLAS Collab.)
ANTIPOV	04a	PL B604 22	Yu.M. Antipov <i>et al.</i>	(IHEP SPHINX Collab.)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HEL5, CIT, CERN)
BARBER	80D	ZPHY C7 17	D.P. Barber <i>et al.</i>	(DARE, LANC, SHEF)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
BARLAG	79	NP B149 220	S.J.M. Barlag <i>et al.</i>	(AMST, CERN, NUM+)
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
CAMERON	77	NP B131 399	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MAST	76	PR D14 13	T.S. Mast <i>et al.</i>	(LBL)
CORDEN	75	NP B84 306	M.J. Corden <i>et al.</i>	(BIRM)
BERTHON	74	NC 21A 146	A. Berthon <i>et al.</i>	(CDEF, RHEL, SACL+)
MAST	73	PR D7 3212	T.S. Mast <i>et al.</i>	(LBL) IJP
MAST	73b	PR D7 5	T.S. Mast <i>et al.</i>	(LBL) IJP
CHAN	72	PRL 28 256	S.B. Chan <i>et al.</i>	(MASA, YALE)
BURKHARDT	71	NP B27 64	E. Burkhardt <i>et al.</i>	(HEID, CERN, SACL)
KIM	71	PRL 27 356	J.K. Kim	(HARV) IJP
Also		Duke Conf. 161	J.K. Kim	(HARV) IJP
Hyperon Resonances, 1970				
BARBARO...	69b	Lund Conf. 352	A. Barbaro-Galteri <i>et al.</i>	(LRL)
Also		Duke Conf. 95	R.D. Tripp	(LRL)
Hyperon Resonances 1970				
BURKHARDT	69	NP B14 106	E. Burkhardt <i>et al.</i>	(HEID, EFI, CERN+)
MAST	68b	PRL 21 1715	T.S. Mast <i>et al.</i>	(LRL)
SCHUEER	68	NP B8 503	J.C. Scheuer <i>et al.</i>	(SABRE Collab.)
DAHL	67	PR 163 1377	O.I. Dahl <i>et al.</i>	(LRL)
DAUBER	67	PL 24B 525	P.M. Dauber <i>et al.</i>	(UCLA)
UHLIG	67	PR 155 1448	R.P. Uhlig <i>et al.</i>	(UMD, NRL)
BIRMINGHAM	66	PR 152 1148	M. Haque <i>et al.</i>	(BIRM, GLAS, LOIC, OXF+)
ARMENTEROS 65C		PL 19 338	R. Armenteros <i>et al.</i>	(CERN, HEID, SACL)
MUSGRAVE 65		NC 35 735	B. Musgrave <i>et al.</i>	(BIRM, CERN, EPOL+)
WATSON	63	PR 131 2248	M.B. Watson, M. Ferro-Luzzi, R.D. Tripp	(BIRM, CERN, EPOL+) IJP
FERRO-LUZZI	62	PRL 8 28	M. Ferro-Luzzi, R.D. Tripp, M.B. Watson	(LRL) IJP

$\Lambda(1600) 1/2^+$

$I(J^P) = 0(\frac{1}{2}^+)$  Status: \*\*\*

See also the  $\Lambda(1810) P_{01}$ . There are quite possibly two  $P_{01}$  states in this region.

$\Lambda(1600)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1560 to 1700 (<math>\approx 1600</math>) OUR ESTIMATE</b>				
1592 $\pm$ 10	ZHANG	13A	DPWA	Multichannel
1568 $\pm$ 20	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1703 $\pm$ 100	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1573 $\pm$ 25	GOPAL	77	DPWA	$\overline{K}N$ multichannel
1596 $\pm$ 6	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
1620 $\pm$ 10	LANGBEIN	72	IPWA	$\overline{K}N$ multichannel
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1572 or 1617	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
1646 $\pm$ 7	<sup>2</sup> CARROLL	76	DPWA	Isospin-0 total $\sigma$
1570	KIM	71	DPWA	K-matrix analysis

$\Lambda(1600)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
50 to 250 ( $\approx 150$ ) OUR ESTIMATE				
150 $\pm$ 28	ZHANG	13A	DPWA	Multichannel
116 $\pm$ 20	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
593 $\pm$ 200	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
147 $\pm$ 50	GOPAL	77	DPWA	$\overline{K}N$ multichannel
175 $\pm$ 20	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
60 $\pm$ 10	LANGBEIN	72	IPWA	$\overline{K}N$ multichannel
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
247 or 271	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
20	<sup>2</sup> CARROLL	76	DPWA	Isospin-0 total $\sigma$
50	KIM	71	DPWA	K-matrix analysis

$\Lambda(1600)$ POLE POSITION				
REAL PART				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1572	ZHANG	13A	DPWA	Multichannel
−2×IMAGINARY PART				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
138	ZHANG	13A	DPWA	Multichannel

$\Lambda(1600)$ DECAY MODES				
Mode	Fraction ( $\Gamma_i/\Gamma$ )			
$\Gamma_1$ $N\bar{K}$	15–30 %			
$\Gamma_2$ $\Sigma\pi$	10–60 %			

The above branching fractions are our estimates, not fits or averages.

$\Lambda(1600)$ BRANCHING RATIOS				
See "Sign conventions for resonance couplings" in the Note on $\Lambda$ and $\Sigma$ Resonances.				
$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.15 to 0.30 OUR ESTIMATE</b>				
0.14 $\pm$ 0.04	ZHANG	13a	DPWA	Multichannel
0.23 $\pm$ 0.04	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.14 $\pm$ 0.05	ALSTON---	78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.25 $\pm$ 0.15	LANGBEIN	72	IPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.24 $\pm$ 0.04	GOPAL	77	DPWA	See GOPAL 80
0.30 or 0.29	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel
$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1600) \rightarrow \Sigma\pi$ <span style="float:right"><math>(\Gamma_1\Gamma_2)^{1/2}/\Gamma</math></span>				
VALUE	DOCUMENT ID	TECN	COMMENT	
− 0.23 $\pm$ 0.03	ZHANG	13a	DPWA	Multichannel
− 0.16 $\pm$ 0.04	GOPAL	77	DPWA	$\bar{K}N$ multichannel
− 0.33 $\pm$ 0.11	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
0.28 $\pm$ 0.09	LANGBEIN	72	IPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
− 0.39 or − 0.39	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel
not seen	HEPP	76b	DPWA	$K^-N \rightarrow \Sigma\pi$

See key on page 601

## Baryon Particle Listings

 $\Lambda(1600)$ ,  $\Lambda(1670)$  $\Lambda(1600)$  FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> A total cross-section bump with  $(J+1/2) \Gamma_{\text{el}} / \Gamma_{\text{total}} = 0.04$ .

 $\Lambda(1600)$  REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL) I
HEPP	76B	PL 65B 487	V. Hepp <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP
LANGBEIN	72	NP B47 477	W. Langbein, F. Wagner	(MPIM) IJP
KIM	71	PRL 27 356	J.K. Kim	(HARV) IJP

 $\Lambda(1670)$  1/2<sup>-</sup>

$$I(J^P) = 0(\frac{1}{2}^-) \text{ Status: } ****$$

The measurements of the mass, width, and elasticity published before 1974 are now obsolete and have been omitted. They were last listed in our 1982 edition Physics Letters **111B** 1 (1982).

 $\Lambda(1670)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1660 to 1680 (<math>\approx 1670</math>) OUR ESTIMATE</b>			
1672 $\pm 3$	ZHANG	13A	DPWA Multichannel
1677.5 $\pm 0.8$	<sup>1</sup> GARCIA-REC...	03	DPWA $\bar{K}N$ multichannel
1673 $\pm 2$	MANLEY	02	DPWA $\bar{K}N$ multichannel
1670.8 $\pm 1.7$	KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$
1667 $\pm 5$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
1671 $\pm 3$	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
1670 $\pm 5$	GOPAL	77	DPWA $\bar{K}N$ multichannel
1675 $\pm 2$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$
1679 $\pm 1$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
1665 $\pm 5$	PREVOST	74	DPWA $K^-N \rightarrow \Sigma(1385)\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1668.9 $\pm 2.0$	ABAEV	96	DPWA $K^-p \rightarrow \Lambda\eta$
1664	<sup>2</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 $\Lambda(1670)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>25 to 50 (<math>\approx 35</math>) OUR ESTIMATE</b>			
29 $\pm 5$	ZHANG	13A	DPWA Multichannel
29.2 $\pm 1.4$	<sup>1</sup> GARCIA-REC...	03	DPWA $\bar{K}N$ multichannel
23 $\pm 6$	MANLEY	02	DPWA $\bar{K}N$ multichannel
34.1 $\pm 3.7$	KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$
29 $\pm 5$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
29 $\pm 5$	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
45 $\pm 10$	GOPAL	77	DPWA $\bar{K}N$ multichannel
46 $\pm 5$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$
40 $\pm 3$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
19 $\pm 5$	PREVOST	74	DPWA $K^-N \rightarrow \Sigma(1385)\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
21.1 $\pm 3.6$	ABAEV	96	DPWA $K^-p \rightarrow \Lambda\eta$
12	<sup>2</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 $\Lambda(1670)$  POLE POSITIONS

## REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1667	ZHANG	13A	DPWA Multichannel

-2 $\times$ IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
26	ZHANG	13A	DPWA Multichannel

 $\Lambda(1670)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	20–30 %
$\Gamma_2$ $\Sigma\pi$	25–55 %
$\Gamma_3$ $\Lambda\eta$	10–25 %
$\Gamma_4$ $\Sigma(1385)\pi$	
The above branching fractions are our estimates, not fits or averages.	
$\Gamma_5$ $N\bar{K}^*(892)$ , $S=3/2$ , $D$ -wave	( $\pm 4$ ) %

 $\Lambda(1670)$  BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

 $\Gamma(N\bar{K})/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.20 to 0.30 OUR ESTIMATE</b>			
0.26 $\pm 0.25$	ZHANG	13A	DPWA Multichannel
0.37 $\pm 0.07$	MANLEY	02	DPWA $\bar{K}N$ multichannel
0.18 $\pm 0.03$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
0.17 $\pm 0.03$	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.20 $\pm 0.03$	GOPAL	77	DPWA See GOPAL 80
0.15	<sup>2</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 $\Gamma(\Lambda\eta)/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.30 $\pm 0.08$	ABAEV	96	DPWA $K^-p \rightarrow \Lambda\eta$

 $(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Lambda(1670) \rightarrow \Sigma\pi$ 

VALUE	DOCUMENT ID	TECN	COMMENT
-0.29 $\pm 0.06$	ZHANG	13A	DPWA Multichannel
-0.38 $\pm 0.03$	MANLEY	02	DPWA $\bar{K}N$ multichannel
-0.26 $\pm 0.02$	KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$
-0.31 $\pm 0.03$	GOPAL	77	DPWA $\bar{K}N$ multichannel
-0.29 $\pm 0.03$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$
-0.23 $\pm 0.03$	LONDON	75	HLBC $K^-p \rightarrow \Sigma^0\pi^0$
-0.27 $\pm 0.02$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.13	<sup>2</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 $(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Lambda(1670) \rightarrow \Lambda\eta$ 

VALUE	DOCUMENT ID	TECN	COMMENT
-0.30 $\pm 0.10$	ZHANG	13A	DPWA Multichannel
+0.24 $\pm 0.04$	MANLEY	02	DPWA $\bar{K}N$ multichannel
+0.20 $\pm 0.05$	BAXTER	73	DPWA $K^-p \rightarrow$ neutrals
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.24	KIM	71	DPWA K-matrix analysis
0.26	ARMENTEROS69C	HBC	
0.20 or 0.23	BERLEY	65	HBC

 $(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Lambda(1670) \rightarrow \Sigma(1385)\pi$ 

VALUE	DOCUMENT ID	TECN	COMMENT
-0.17 $\pm 0.06$	MANLEY	02	DPWA $\bar{K}N$ multichannel
-0.18 $\pm 0.05$	PREVOST	74	DPWA $K^-N \rightarrow \Sigma(1385)\pi$

 $\Gamma(N\bar{K}^*(892), S=3/2, D\text{-wave})/\Gamma_{\text{total}}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.05 <math>\pm 0.04</math></b>	ZHANG	13A	DPWA Multichannel

 $\Lambda(1670)$  FOOTNOTES

- <sup>1</sup> GARCIA-RECIO 03 gives pole, not Breit-Wigner, parameters, but the narrow width of the  $\Lambda(1670)$  means there will be little difference.  
<sup>2</sup> MARTIN 77 obtains identical resonance parameters from a T-matrix pole and from a Breit-Wigner fit.

 $\Lambda(1670)$  REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GARCIA-REC...	03	PR D67 076009	C. Garcia-Recio <i>et al.</i>	(GRAN, VALE)
MANLEY	02	PRL 88 012002	D.M. Manley <i>et al.</i>	(BNL Crystal Ball Collab.)
ABAEV	96	PR C53 385	V.V. Aboev, B.M.K. Nefkens	(UCLA)
KOISO	85	NP A433 619	H. Koiso <i>et al.</i>	(TOKYO, MASA)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HELS, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
HEPP	76B	PL 65B 487	V. Hepp <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
LONDON	75	NP B85 289	G.W. London <i>et al.</i>	(BNL, CERN, EPOL+) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i>	(SACL, CERN, HEID)
BAXTER	73	NP B67 125	D.F. Baxter <i>et al.</i>	(OXF) IJP
KIM	71	PRL 27 356	J.K. Kim	(HARV) IJP
Also		Duke Conf. 161	J.K. Kim	(HARV) IJP
Hyperon Resonances, 1970				
ARMENTEROS 69C		Lund Paper 229	R. Armenteros <i>et al.</i>	(CERN, HEID, SACL) IJP
Values are quoted in LEVI-SETTI 69.				
BERLEY	65	PRL 15 641	D. Berley <i>et al.</i>	(BNL) IJP

Baryon Particle Listings

$\Lambda(1690), \Lambda(1710)$

$\Lambda(1690) \ 3/2^-$

$I(J^P) = 0(\frac{3}{2}^-)$  Status: \*\*\*\*

The measurements of the mass, width, and elasticity published before 1974 are now obsolete and have been omitted. They were last listed in our 1982 edition Physics Letters **111B** 1 (1982).

$\Lambda(1690)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1685 to 1695 (<math>\approx 1690</math>) OUR ESTIMATE</b>			
1691 $\pm 3$	ZHANG	13A	DPWA Multichannel
1695.7 $\pm 2.6$	KOISO	85	DPWA $K^-p \rightarrow \Sigma \pi$
1690 $\pm 5$	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$
1692 $\pm 5$	ALSTON...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$
1690 $\pm 5$	GOPAL	77	DPWA $\overline{K}N$ multichannel
1690 $\pm 3$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma \pi$
1689 $\pm 1$	KANE	74	DPWA $K^-p \rightarrow \Sigma \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1687 or 1689	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
1692 $\pm 4$	CARROLL	76	DPWA Isospin-0 total $\sigma$

$\Lambda(1690)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>50 to 70 (<math>\approx 60</math>) OUR ESTIMATE</b>			
54 $\pm 5$	ZHANG	13A	DPWA Multichannel
67.2 $\pm 5.6$	KOISO	85	DPWA $K^-p \rightarrow \Sigma \pi$
61 $\pm 5$	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$
64 $\pm 10$	ALSTON...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$
60 $\pm 5$	GOPAL	77	DPWA $\overline{K}N$ multichannel
82 $\pm 8$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma \pi$
60 $\pm 4$	KANE	74	DPWA $K^-p \rightarrow \Sigma \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
62 or 62	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
38	CARROLL	76	DPWA Isospin-0 total $\sigma$

$\Lambda(1690)$ POLE POSITION			
<b>REAL PART</b>			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1689	ZHANG	13A	DPWA Multichannel
<b>-2xIMAGINARY PART</b>			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
53	ZHANG	13A	DPWA Multichannel

$\Lambda(1690)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \ N\overline{K}$	20-30 %
$\Gamma_2 \ \Sigma \pi$	20-40 %
$\Gamma_3 \ \Lambda \pi \pi$	$\sim 25$ %
$\Gamma_4 \ \Sigma \pi \pi$	$\sim 20$ %
$\Gamma_5 \ \Lambda \eta$	
$\Gamma_6 \ \Sigma(1385) \pi, S\text{-wave}$	
The above branching fractions are our estimates, not fits or averages.	

$\Lambda(1690)$ BRANCHING RATIOS			
The sum of all the quoted branching ratios is more than 1.0. The two-body ratios are from partial-wave analyses, and thus probably are more reliable than the three-body ratios, which are determined from bumps in cross sections. Of the latter, the $\Sigma \pi \pi$ bump looks more significant. (The error given for the $\Lambda \pi \pi$ ratio looks unreasonably small.) Hardly any of the $\Sigma \pi \pi$ decay can be via $\Sigma(1385)$ , for then seven times as much $\Lambda \pi \pi$ decay would be required. See "Sign conventions for resonance couplings" in the Note on $\Lambda$ and $\Sigma$ Resonances.			
$\Gamma(N\overline{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
<b>0.2 to 0.3 OUR ESTIMATE</b>			
0.25 $\pm 0.04$	ZHANG	13A	DPWA Multichannel
0.23 $\pm 0.03$	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$
0.22 $\pm 0.03$	ALSTON...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.24 $\pm 0.03$	GOPAL	77	DPWA See GOPAL 80
0.28 or 0.26	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Lambda(1690) \rightarrow \Sigma \pi$			
VALUE	DOCUMENT ID	TECN	COMMENT
-0.27 $\pm 0.03$	ZHANG	13A	DPWA Multichannel
-0.34 $\pm 0.02$	KOISO	85	DPWA $K^-p \rightarrow \Sigma \pi$
-0.25 $\pm 0.03$	GOPAL	77	DPWA $\overline{K}N$ multichannel
-0.29 $\pm 0.03$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma \pi$
-0.28 $\pm 0.03$	LONDON	75	HLBC $K^-p \rightarrow \Sigma^0 \pi^0$
-0.28 $\pm 0.02$	KANE	74	DPWA $K^-p \rightarrow \Sigma \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.30 or -0.28	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Lambda(1690) \rightarrow \Lambda \pi \pi$			
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.25 $\pm 0.02$	<sup>2</sup> BARTLEY	68	HDBC $K^-p \rightarrow \Lambda \pi \pi$

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Lambda(1690) \rightarrow \Sigma \pi \pi$			
VALUE	DOCUMENT ID	TECN	COMMENT
0.21	ARMENTEROS68c	HDBC	$K^-N \rightarrow \Sigma \pi \pi$

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Lambda(1690) \rightarrow \Lambda \eta$			
VALUE	DOCUMENT ID	TECN	COMMENT
0.00 $\pm 0.03$	BAXTER	73	DPWA $K^-p \rightarrow$ neutrals

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Lambda(1690) \rightarrow \Sigma(1385) \pi, S\text{-wave}$			
VALUE	DOCUMENT ID	TECN	COMMENT
-0.28 $\pm 0.06$	ZHANG	13A	DPWA Multichannel
+0.27 $\pm 0.04$	PREVOST	74	DPWA $K^-N \rightarrow \Sigma(1385) \pi$

- $\Lambda(1690)$  FOOTNOTES**
- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit. Another  $D_{03} \ \Lambda$  at 1966 MeV is also suggested by MARTIN 77, but is very uncertain.

<sup>2</sup> BARTLEY 68 uses only cross-section data. The enhancement is not seen by PREVOST 71.

$\Lambda(1690)$ REFERENCES			
ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i> (KSU)
KOISO	85	NP A433 619	H. Koiso <i>et al.</i> (TOKY, MASA)
PDG	82	PL 111B 1	M. Roos <i>et al.</i> (HELS, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal (RHEL) IJP
ALSTON...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i> (LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i> (LBL, MTHO+) IJP
GOPAL	77	NP B119 322	G.P. Gopal <i>et al.</i> (LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse (LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pidcock (LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Pidcock (LOUC) IJP
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i> (BNL) I
HEPP	76B	PL 65B 487	V. Hepp <i>et al.</i> (CERN, HEIDH, MPIM) IJP
LONDON	75	NP B85 289	G.W. London <i>et al.</i> (BNL, CERN, EPOL+) IJP
KANE	74	LBL-2452	D.F. Kane (LBL) IJP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i> (SACL, CERN, HEID) IJP
BAXTER	73	NP B67 125	D.F. Baxter <i>et al.</i> (OXF) IJP
PREVOST	71	Amsterdam Conf.	J. Prevost (CERN, HEID, SACL) IJP
ARMENTEROS	68C	NP B8 216	R. Armenteros <i>et al.</i> (CERN, HEID, SACL) I
BARTLEY	68	PRL 21 1111	J.H. Bartley <i>et al.</i> (TUFTS, FSU, BRAN) I

$\Lambda(1710) \ 1/2^+$

$I(J^P) = 0(\frac{1}{2}^+)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Lambda(1710)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1713 <math>\pm 13</math></b>	ZHANG	13A	DPWA Multichannel

$\Lambda(1710)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>180 <math>\pm 42</math></b>	ZHANG	13A	DPWA Multichannel

$\Lambda(1710)$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \ N\overline{K}$	(43 $\pm 4$ ) %
$\Gamma_2 \ \Sigma \pi$	(21 $\pm 5$ ) %
$\Gamma_3 \ \Sigma^*(1385) \pi, P\text{-wave}$	(20 $\pm 8$ ) %
$\Gamma_4 \ N\overline{K}^*(892)$	
$\Gamma_5 \ N\overline{K}^*(892), S=1/2$	( 5 $\pm 4$ ) %
$\Gamma_6 \ N\overline{K}^*(892), S=3/2, P\text{-wave}$	(10 $\pm 8$ ) %



See key on page 601

Baryon Particle Listings  
 $\Lambda(1710)$ ,  $\Lambda(1800)$

$\Lambda(1710)$ BRANCHING RATIOS				
$\Gamma(N\bar{K})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.43 \pm 0.04</math></b>	ZHA NG	13A	DPWA	Multichannel
$\Gamma(\Sigma\pi)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.21 \pm 0.05</math></b>	ZHA NG	13A	DPWA	Multichannel
$\Gamma(\Sigma^*(1385)\pi, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.20 \pm 0.08</math></b>	ZHA NG	13A	DPWA	Multichannel
$\Gamma(N\bar{K}^*(892), S=1/2)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.05 \pm 0.04</math></b>	ZHA NG	13A	DPWA	Multichannel
$\Gamma(N\bar{K}^*(892), S=3/2, P\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.10 \pm 0.08</math></b>	ZHA NG	13A	DPWA	Multichannel

$\Lambda(1710)$ REFERENCES				
ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
<div><div><math>\Lambda(1800) 1/2^-</math></div><div><math>I(J^P) = 0(\frac{1}{2}^-)</math> Status: ***</div></div> <p>This is the second resonance in the <math>S_{01}</math> wave, the first being the <math>\Lambda(1670)</math>.</p>				

$\Lambda(1800)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1720 to 1850 (<math>\approx 1800</math>) OUR ESTIMATE</b>				
1783 $\pm$ 19	ZHA NG	13A	DPWA	Multichannel
1845 $\pm$ 10	MANLEY	02	DPWA	$\bar{K}N$ multichannel
1841 $\pm$ 10	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
1725 $\pm$ 20	ALSTON-...	78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
1825 $\pm$ 20	GOPAL	77	DPWA	$\bar{K}N$ multichannel
1830 $\pm$ 20	LANGBEIN	72	IPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1767 or 1842	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel
1780	KIM	71	DPWA	K-matrix analysis
1872 $\pm$ 10	BRICMAN	70B	DPWA	$\bar{K}N \rightarrow \bar{K}N$

$\Lambda(1800)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>200 to 400 (<math>\approx 300</math>) OUR ESTIMATE</b>				
256 $\pm$ 35	ZHA NG	13A	DPWA	Multichannel
518 $\pm$ 84	MANLEY	02	DPWA	$\bar{K}N$ multichannel
228 $\pm$ 20	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
185 $\pm$ 20	ALSTON-...	78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
230 $\pm$ 20	GOPAL	77	DPWA	$\bar{K}N$ multichannel
70 $\pm$ 15	LANGBEIN	72	IPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
435 or 473	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel
40	KIM	71	DPWA	K-matrix analysis
100 $\pm$ 20	BRICMAN	70B	DPWA	$\bar{K}N \rightarrow \bar{K}N$

$\Lambda(1800)$ POLE POSITION				
REAL PART				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1729	ZHA NG	13A	DPWA	Multichannel
$-2 \times$ IMAGINARY PART				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
198	ZHA NG	13A	DPWA	Multichannel

$\Lambda(1800)$ DECAY MODES		
	Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$N\bar{K}$	25–40 %
$\Gamma_2$	$\Sigma\pi$	seen
$\Gamma_3$	$\Sigma(1385)\pi$	seen
	The above branching fractions are our estimates, not fits or averages.	
$\Gamma_4$	$\Lambda\eta$	(6 $\pm$ 5) %
$\Gamma_5$	$N\bar{K}^*(892)$	seen
$\Gamma_6$	$N\bar{K}^*(892)$ , $S=1/2$ , $S$ -wave	
$\Gamma_7$	$N\bar{K}^*(892)$ , $S=3/2$ , $D$ -wave	

$\Lambda(1800)$ BRANCHING RATIOS				
See “Sign conventions for resonance couplings” in the Note on $\Lambda$ and $\Sigma$ Resonances.				
$\Gamma(N\bar{K})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.25 to 0.40 OUR ESTIMATE</b>				
0.13 $\pm$ 0.06	ZHANG	13A	DPWA	Multichannel
0.24 $\pm$ 0.10	MANLEY	02	DPWA	$\bar{K}N$ multichannel
0.36 $\pm$ 0.04	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.28 $\pm$ 0.05	ALSTON-...	78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.35 $\pm$ 0.15	LANGBEIN	72	IPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.37 $\pm$ 0.05	GOPAL	77	DPWA	See GOPAL 80
1.21 or 0.70	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel
0.80	KIM	71	DPWA	K-matrix analysis
0.18 $\pm$ 0.02	BRICMAN	70B	DPWA	$\bar{K}N \rightarrow \bar{K}N$

$(\Gamma_1\Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1800) \rightarrow \Sigma\pi$				$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
– 0.07 $\pm$ 0.02	ZHANG	13A	DPWA	Multichannel
– 0.08 $\pm$ 0.05	GOPAL	77	DPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
– 0.74 or – 0.43	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel
0.24	KIM	71	DPWA	K-matrix analysis

$(\Gamma_1\Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1800) \rightarrow \Sigma(1385)\pi$				$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
– 0.09 $\pm$ 0.05	ZHANG	13A	DPWA	Multichannel
+ 0.056 $\pm$ 0.028	<sup>2</sup> CAMERON	78	DPWA	$K^-p \rightarrow \Sigma(1385)\pi$

$\Gamma(\Lambda\eta)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>0.06 \pm 0.05</math></b>	ZHANG	13A	DPWA	Multichannel

$(\Gamma_1\Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1800) \rightarrow N\bar{K}^*(892), S=1/2, S\text{-wave}$				$(\Gamma_1\Gamma_6)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
– 0.13 $\pm$ 0.02	ZHANG	13A	DPWA	Multichannel
– 0.17 $\pm$ 0.03	<sup>2</sup> CAMERON	78B	DPWA	$K^-p \rightarrow N\bar{K}^*$

$(\Gamma_1\Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1800) \rightarrow N\bar{K}^*(892), S=3/2, D\text{-wave}$				$(\Gamma_1\Gamma_7)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
– 0.13 $\pm$ 0.04	CAMERON	78B	DPWA	$K^-p \rightarrow N\bar{K}^*$

$\Lambda(1800)$ FOOTNOTES	
<sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.	
<sup>2</sup> The published sign has been changed to be in accord with the baryon-first convention.	

$\Lambda(1800)$ REFERENCES				
ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
MANLEY	02	PRL 88 012002	D.M. Manley <i>et al.</i>	(BNL Crystal Ball Collab.)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also	78	PR 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
CAMERON	78B	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
GOPAL	77	NP B119 362	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
MARTIN	77	NP B127 349	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
Also	77	NP B126 266	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also	77	NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
LANGBEIN	72	NP B47 477	B.R. Martin, M.K. Pidcock	(MPIM) IJP
KIM	71	PRL 27 356	W. Langbein, F. Wagner	(HARV) IJP
Also	71	Duke Conf. 161	J.K. Kim	(HARV) IJP
Hyperon Resonances, 1970			J.K. Kim	(HARV) IJP
BRICMAN	70B	PL 33B 511	C. Bricean, M. Ferro-Luzzi, J.P. Lagnaux	(CERN) IJP

# Baryon Particle Listings

## $\Lambda(1810), \Lambda(1820)$

$\Lambda(1810) \ 1/2^+$

$I(J^P) = 0(\frac{1}{2}^+)$  Status: \*\*\*

Almost all the recent analyses contain a  $P_{01}$  state, and sometimes two of them, but the masses, widths, and branching ratios vary greatly. See also the  $\Lambda(1600)$   $P_{01}$ .

$\Lambda(1810)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1750 to 1850 (<math>\approx 1810</math>) OUR ESTIMATE</b>				
1821 $\pm$ 10	ZHANG	13A	DPWA	Multichannel
1841 $\pm$ 20	GOPAL	80	DPWA	$\overline{K} N \rightarrow \overline{K} N$
1853 $\pm$ 20	GOPAL	77	DPWA	$\overline{K} N$ multichannel
1735 $\pm$ 5	CARROLL	76	DPWA	Isospin-0 total $\sigma$
1746 $\pm$ 10	PREVOST	74	DPWA	$K^- N \rightarrow \Sigma(1385) \pi$
1780 $\pm$ 20	LANGBEIN	72	IPWA	$\overline{K} N$ multichannel
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1861 or 1953	<sup>1</sup> MARTIN	77	DPWA	$\overline{K} N$ multichannel
1755	KIM	71	DPWA	K-matrix analysis
1800	ARMENTEROS70	HBC	$\overline{K} N \rightarrow \overline{K} N$	
1750	ARMENTEROS70	HBC	$\overline{K} N \rightarrow \Sigma \pi$	
1690 $\pm$ 10	BARBARO-...	70	HBC	$\overline{K} N \rightarrow \Sigma \pi$
1740	BAILEY	69	DPWA	$\overline{K} N \rightarrow \overline{K} N$
1745	ARMENTEROS68B	HBC	$\overline{K} N \rightarrow \overline{K} N$	

$\Lambda(1810)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>50 to 250 (<math>\approx 150</math>) OUR ESTIMATE</b>				
174 $\pm$ 5	ZHANG	13A	DPWA	Multichannel
164 $\pm$ 20	GOPAL	80	DPWA	$\overline{K} N \rightarrow \overline{K} N$
90 $\pm$ 20	CAMERON	78B	DPWA	$K^- p \rightarrow N \overline{K}^*$
166 $\pm$ 20	GOPAL	77	DPWA	$\overline{K} N$ multichannel
46 $\pm$ 20	PREVOST	74	DPWA	$K^- N \rightarrow \Sigma(1385) \pi$
120 $\pm$ 10	LANGBEIN	72	IPWA	$\overline{K} N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
535 or 585	<sup>1</sup> MARTIN	77	DPWA	$\overline{K} N$ multichannel
28	CARROLL	76	DPWA	Isospin-0 total $\sigma$
35	KIM	71	DPWA	K-matrix analysis
30	ARMENTEROS70	HBC	$\overline{K} N \rightarrow \overline{K} N$	
70	ARMENTEROS70	HBC	$\overline{K} N \rightarrow \Sigma \pi$	
22	BARBARO-...	70	HBC	$\overline{K} N \rightarrow \Sigma \pi$
300	BAILEY	69	DPWA	$\overline{K} N \rightarrow \overline{K} N$
147	ARMENTEROS68B	HBC		

$\Lambda(1810)$ POLE POSITION				
<b>REAL PART</b>				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1780	ZHA NG	13A	DPWA	Multichannel
<b>−2×IMAGINARY PART</b>				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
64	ZHA NG	13A	DPWA	Multichannel

$\Lambda(1810)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$	$N\overline{K}$	20–50 %
$\Gamma_2$	$\Sigma \pi$	10–40 %
$\Gamma_3$	$\Sigma(1385)\pi$	seen
$\Gamma_4$	$N\overline{K}^*(892)$	30–60 %
$\Gamma_5$	$N\overline{K}^*(892)$ , $S=1/2$ , $P$ -wave	
$\Gamma_6$	$N\overline{K}^*(892)$ , $S=3/2$ , $P$ -wave	
The above branching fractions are our estimates, not fits or averages.		

$\Lambda(1810)$ BRANCHING RATIOS				
See “Sign conventions for resonance couplings” in the Note on $\Lambda$ and $\Sigma$ Resonances.				
$\Gamma(N\overline{K})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.2 to 0.5 OUR ESTIMATE</b>				
0.19±0.08	ZHANG	13a	DPWA	Multichannel
0.24±0.04	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
0.36±0.05	LANGBEIN	72	IPWA	$\overline{K}N$ multichannel

• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.21 $\pm$ 0.04	GOPAL	77	DPWA	See GOPAL 80
0.52 or 0.49	<sup>1</sup> MARTIN	77	DPWA	$\overline{K} N$ multichannel
0.30	KIM	71	DPWA	K-matrix analysis
0.15	ARMENTEROS70		DPWA	$\overline{K} N \rightarrow \overline{K} N$
0.55	BAILEY	69	DPWA	$\overline{K} N \rightarrow \overline{K} N$
0.4	ARMENTEROS68B		DPWA	$\overline{K} N \rightarrow \overline{K} N$

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N \overline{K} \rightarrow \Lambda(1810) \rightarrow \Sigma \pi$				
VALUE	DOCUMENT ID	TECN	COMMENT	
-0.08 $\pm$ 0.05	ZHANG	13A	DPWA	Multichannel
-0.24 $\pm$ 0.04	GOPAL	77	DPWA	$\overline{K} N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
+0.25 or +0.23	<sup>1</sup> MARTIN	77	DPWA	$\overline{K} N$ multichannel
< 0.01	LANGBEIN	72	IPWA	$\overline{K} N$ multichannel
0.17	KIM	71	DPWA	K-matrix analysis
+0.20	<sup>2</sup> ARMENTEROS70		DPWA	$\overline{K} N \rightarrow \Sigma \pi$
-0.13 $\pm$ 0.03	BARBARO-...	70	DPWA	$\overline{K} N \rightarrow \Sigma \pi$

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N \overline{K} \rightarrow \Lambda(1810) \rightarrow \Sigma(1385) \pi$				
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.18 $\pm$ 0.10	PREVOST	74	DPWA	$K^- N \rightarrow \Sigma(1385) \pi$

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N \overline{K} \rightarrow \Lambda(1810) \rightarrow N \overline{K}^*(892), S=1/2, P\text{-wave}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
-0.14 $\pm$ 0.03	<sup>2</sup> CAMERON	78B	DPWA	$K^- p \rightarrow N \overline{K}^*$

$(\Gamma_i \Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N \overline{K} \rightarrow \Lambda(1810) \rightarrow N \overline{K}^*(892), S=3/2, P\text{-wave}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.38 $\pm$ 0.06	ZHANG	13A	DPWA	Multichannel
+0.35 $\pm$ 0.06	CAMERON	78B	DPWA	$K^- p \rightarrow N \overline{K}^*$

$\Lambda(1810)$ FOOTNOTES	
<sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.	
<sup>2</sup> The published sign has been changed to be in accord with the baryon-first convention.	

$\Lambda(1810)$ REFERENCES			
ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i> (KSU)
GOPAL	80	Toronto Conf. 159	G.P. Gopal (RHEL) IJP
CAMERON	78B	NP B146 327	W. Cameron <i>et al.</i> (RHEL, LOIC) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i> (LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse (LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pidcock (LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Pidcock (BNL) I
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i> (SACL, CERN, HEID) I
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i> (MPIM) IJP
LANGBEIN	72	NP B47 477	W. Langbein, F. Wagner (HARV) IJP
KIM	71	PRL 27 356	J.K. Kim (HARV) IJP
Also		Duke Conf. 161	
Hyperon Resonances, 1970			
ARMENTEROS 70		Duke Conf. 123	R. Armenteros <i>et al.</i> (CERN, HEID, SACL) IJP
Hyperon Resonances, 1970			
BARBARO-...	70	Duke Conf. 173	A. Barbaro-Galtieri (LRL) IJP
Hyperon Resonances, 1970			
BAILEY	69	Thesis UCRL 50617	J.M. Bailey (LLL) IJP
ARMENTEROS 68B		NP B8 195	R. Armenteros <i>et al.</i> (CERN, HEID, SACL) IJP

$\Lambda(1820) \ 5/2^+$

$I(J^P) = 0(\frac{5}{2}^+)$  Status: \*\*\*

This resonance is the cornerstone for all partial-wave analyses in this region. Most of the results published before 1973 are now obsolete and have been omitted. They may be found in our 1982 edition Physics Letters **111B** 1 (1982).

Most of the quoted errors are statistical only; the systematic errors due to the particular parametrizations used in the partial-wave analyses are not included. For this reason we do not calculate weighted averages for the mass and width.

$\Lambda(1820)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1815 to 1825 (<math>\approx 1820</math>) OUR ESTIMATE</b>				
1823.5 $\pm$ 0.8	ZHANG	13A	DPWA	Multichannel
1823 $\pm$ 3	GOPAL	80	DPWA	$\overline{K} N \rightarrow \overline{K} N$
1819 $\pm$ 2	ALSTON-...	78	DPWA	$\overline{K} N \rightarrow \overline{K} N$
1822 $\pm$ 2	GOPAL	77	DPWA	$\overline{K} N$ multichannel
1821 $\pm$ 2	KANE	74	DPWA	$K^- p \rightarrow \Sigma \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1830	DECLAIS	77	DPWA	$\overline{K} N \rightarrow \overline{K} N$
1817 or 1819	<sup>1</sup> MARTIN	77	DPWA	$\overline{K} N$ multichannel

See key on page 601

# Baryon Particle Listings

## $\Lambda(1820)$ , $\Lambda(1830)$

 **$\Lambda(1820)$  WIDTH**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>70 to 90 (<math>\approx 80</math>) OUR ESTIMATE</b>			
$89 \pm 2$	ZHANG	13A	DPWA Multichannel
$77 \pm 5$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
$72 \pm 5$	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
$81 \pm 5$	GOPAL	77	DPWA $\bar{K}N$ multichannel
$87 \pm 3$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
82	DECLAIS	77	DPWA $\bar{K}N \rightarrow \bar{K}N$
76 or 76	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 **$\Lambda(1820)$  POLE POSITION****REAL PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1814	ZHANG	13A	DPWA Multichannel

**-2×IMAGINARY PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
85	ZHANG	13A	DPWA Multichannel

 **$\Lambda(1820)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	55–65 %
$\Gamma_2$ $\Sigma\pi$	8–14 %
$\Gamma_3$ $\Sigma(1385)\pi$	5–10 %
$\Gamma_4$ $\Sigma(1385)\pi, P\text{-wave}$	
$\Gamma_5$ $\Sigma(1385)\pi, F\text{-wave}$	
$\Gamma_6$ $\Lambda\eta$	
$\Gamma_7$ $\Sigma\pi\pi$	
The above branching fractions are our estimates, not fits or averages.	
$\Gamma_8$ $N\bar{K}^*(892), S=3/2, P\text{-wave}$	( $3.0 \pm 1.0$ ) %

 **$\Lambda(1820)$  BRANCHING RATIOS**

Errors quoted do not include uncertainties in the parametrizations used in the partial-wave analyses and are thus too small. See also “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE				
<b>0.55 to 0.65 OUR ESTIMATE</b>				
$0.54 \pm 0.01$	ZHANG	13A	DPWA Multichannel	
$0.58 \pm 0.02$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$	
$0.60 \pm 0.03$	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.51	DECLAIS	77	DPWA $\bar{K}N \rightarrow \bar{K}N$	
$0.57 \pm 0.02$	GOPAL	77	DPWA See GOPAL 80	
0.59 or 0.58	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel	

$(\Gamma_1/\Gamma)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1820) \rightarrow \Sigma\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE				
$-0.28 \pm 0.01$	ZHANG	13A	DPWA Multichannel	
$-0.28 \pm 0.03$	GOPAL	77	DPWA $\bar{K}N$ multichannel	
$-0.28 \pm 0.01$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.25$ or $-0.25$	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel	

$\Gamma(\Sigma\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
VALUE				
no clear signal	<sup>2</sup> ARMENTEROS68C	HDBC	$K^-N \rightarrow \Sigma\pi\pi$	

$(\Gamma_1/\Gamma)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1820) \rightarrow \Sigma(1385)\pi, P\text{-wave}$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$
VALUE				
$-0.20 \pm 0.02$	ZHANG	13A	DPWA Multichannel	
$-0.167 \pm 0.054$	<sup>3</sup> CAMERON	78	DPWA $K^-p \rightarrow \Sigma(1385)\pi$	
$+0.27 \pm 0.03$	PREVOST	74	DPWA $K^-N \rightarrow \Sigma(1385)\pi$	

$(\Gamma_1/\Gamma)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1820) \rightarrow \Sigma(1385)\pi, F\text{-wave}$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$
VALUE				
$+0.065 \pm 0.029$	<sup>3</sup> CAMERON	78	DPWA $K^-p \rightarrow \Sigma(1385)\pi$	

$(\Gamma_1/\Gamma)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1820) \rightarrow \Lambda\eta$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_6)^{1/2}/\Gamma$
VALUE				
$-0.096 \pm 0.040$ $-0.020$	RADER	73	MPWA	

$\Gamma(N\bar{K}^*(892), S=3/2, P\text{-wave})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
VALUE				
<b>0.03 <math>\pm</math> 0.01</b>	ZHANG	13A	DPWA Multichannel	

 **$\Lambda(1820)$  FOOTNOTES**

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> There is a suggestion of a bump, enough to be consistent with what is expected from  $\Sigma(1385) \rightarrow \Sigma\pi$  decay.  
<sup>3</sup> The published sign has been changed to be in accord with the baryon-first convention.

 **$\Lambda(1820)$  REFERENCES**

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HEL5, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
CAMERON	78	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
DECLAIS	77	CERN 77-16	Y. Declais <i>et al.</i>	(CAEN, CERN) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i>	(SACL, CERN, HEID)
RADER	73	NC 16A 178	R.K. Rader <i>et al.</i>	(SACL, HEID, CERN+)
ARMENTEROS	68C	NP B8 216	R. Armenteros <i>et al.</i>	(CERN, HEID, SACL) I

<b><math>\Lambda(1830)</math> 5/2<sup>-</sup></b>	$I(J^P) = 0(\frac{5}{2}^-)$ Status: * * * *
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For results published before 1973 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

The best evidence for this resonance is in the  $\Sigma\pi$  channel.

 **$\Lambda(1830)$  MASS**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1810 to 1830 (<math>\approx 1830</math>) OUR ESTIMATE</b>			
$1820 \pm 4$	ZHANG	13A	DPWA Multichannel
$1831 \pm 10$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
$1825 \pm 10$	GOPAL	77	DPWA $\bar{K}N$ multichannel
$1825 \pm 1$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1817 or 1818	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 **$\Lambda(1830)$  WIDTH**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>60 to 110 (<math>\approx 95</math>) OUR ESTIMATE</b>			
$114 \pm 10$	ZHANG	13A	DPWA Multichannel
$100 \pm 10$	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
$94 \pm 10$	GOPAL	77	DPWA $\bar{K}N$ multichannel
$119 \pm 3$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
56 or 56	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

 **$\Lambda(1830)$  POLE POSITION****REAL PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1809	ZHANG	13A	DPWA Multichannel

**-2×IMAGINARY PART**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
109	ZHANG	13A	DPWA Multichannel

 **$\Lambda(1830)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	3–10 %
$\Gamma_2$ $\Sigma\pi$	35–75 %
$\Gamma_3$ $\Sigma(1385)\pi$	>15 %
The above branching fractions are our estimates, not fits or averages.	
$\Gamma_4$ $\Sigma(1385)\pi, D\text{-wave}$	( $52 \pm 6$ ) %
$\Gamma_5$ $\Lambda\eta$	

# Baryon Particle Listings

## $\Lambda(1830)$ , $\Lambda(1890)$

### $\Lambda(1830)$ BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.03 to 0.10 OUR ESTIMATE</b>				
0.041 ± 0.005	ZHANG	13A	DPWA	Multichannel
0.08 ± 0.03	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.02 ± 0.02	ALSTON-...	78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
0.04 ± 0.03	GOPAL	77	DPWA	See GOPAL 80
0.04 or 0.04	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1830) \rightarrow \Sigma\pi$				$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.13 ± 0.01	ZHANG	13A	DPWA	Multichannel
−0.17 ± 0.03	GOPAL	77	DPWA	$\bar{K}N$ multichannel
−0.15 ± 0.01	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
−0.17 or −0.17	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1830) \rightarrow \Sigma(1385)\pi$				$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.141 ± 0.014	<sup>2</sup> CAMERON	78	DPWA	$K^-p \rightarrow \Sigma(1385)\pi$
+0.13 ± 0.03	PREVOST	74	DPWA	$K^-N \rightarrow \Sigma(1385)\pi$

$\Gamma(\Sigma(1385)\pi, D\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.52 ± 0.06</b>	ZHANG	13A	DPWA	Multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1830) \rightarrow \Lambda\eta$				$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.044 ± 0.020	RADER	73	MPWA	

### $\Lambda(1830)$ FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> The CAMERON 78 upper limit on G-wave decay is 0.03. The published sign has been changed to be in accord with the baryon-first convention.

### $\Lambda(1830)$ REFERENCES

ZHANG	13A	PR C98 035205	H. Zhang <i>et al.</i>	(KSU)
PDG	82	PL 111B.1	M. Roos <i>et al.</i>	(HELSE, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IUP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IUP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IUP
CAMERON	78	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) IUP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IUP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IUP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC) IUP
Also		NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IUP
KANE	74	LBL-2452	D.F. Kane	(LBL) IUP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i>	(SACL, CERN, HEID)
RADER	73	NC 16A 178	R.K. Rader <i>et al.</i>	(SACL, HEID, CERN+)

## $\Lambda(1890)$ 3/2<sup>+</sup>

$I(J^P) = 0(\frac{3}{2}^+)$  Status: \* \* \* \*

For results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

The  $J^P = 3/2^+$  assignment is consistent with all available data (including polarization) and recent partial-wave analyses. The dominant inelastic modes remain unknown.

### $\Lambda(1890)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1850 to 1910 (<math>\approx 1890</math>) OUR ESTIMATE</b>			
1900 ± 5	ZHANG	13A	DPWA
1897 ± 5	GOPAL	80	DPWA
1908 ± 10	ALSTON-...	78	DPWA
1900 ± 5	GOPAL	77	DPWA
1894 ± 10	HEMINGWAY	75	DPWA
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
1856 or 1868	<sup>1</sup> MARTIN	77	DPWA
1900	<sup>2</sup> NAKKASYAN	75	DPWA

### $\Lambda(1890)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>60 to 200 (<math>\approx 100</math>) OUR ESTIMATE</b>			
161 ± 15	ZHANG	13A	DPWA
74 ± 10	GOPAL	80	DPWA
119 ± 20	ALSTON-...	78	DPWA
72 ± 10	GOPAL	77	DPWA
107 ± 10	HEMINGWAY	75	DPWA
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
191 or 193	<sup>1</sup> MARTIN	77	DPWA
100	<sup>2</sup> NAKKASYAN	75	DPWA

### $\Lambda(1890)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
1876	ZHANG	13A	DPWA

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc.	• • •		
145	ZHANG	13A	DPWA

### $\Lambda(1890)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	20–35 %
$\Gamma_2$ $\Sigma\pi$	3–10 %
$\Gamma_3$ $\Sigma(1385)\pi$	seen
$\Gamma_4$ $\Sigma(1385)\pi, P\text{-wave}$	
$\Gamma_5$ $\Sigma(1385)\pi, F\text{-wave}$	
$\Gamma_6$ $N\bar{K}^*(892)$	seen
$\Gamma_7$ $N\bar{K}^*(892), S=1/2$	
$\Gamma_8$ $N\bar{K}^*(892), S=3/2, F\text{-wave}$	
$\Gamma_9$ $\Lambda\omega$	

The above branching fractions are our estimates, not fits or averages.

### $\Lambda(1890)$ BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.20 to 0.35 OUR ESTIMATE</b>				
0.37 ± 0.03	ZHANG	13A	DPWA	Multichannel
0.20 ± 0.02	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.34 ± 0.05	ALSTON-...	78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.24 ± 0.04	HEMINGWAY	75	DPWA	$K^-p \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
0.18 ± 0.02	GOPAL	77	DPWA	See GOPAL 80
0.36 or 0.34	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1890) \rightarrow \Sigma\pi$				$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.09 ± 0.02	ZHANG	13A	DPWA	Multichannel
−0.09 ± 0.03	GOPAL	77	DPWA	$\bar{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc.	• • •			
+0.15 or +0.14	<sup>1</sup> MARTIN	77	DPWA	$\bar{K}N$ multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1890) \rightarrow \Sigma(1385)\pi, P\text{-wave}$				$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
<0.03	CAMERON	78	DPWA	$K^-p \rightarrow \Sigma(1385)\pi$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1890) \rightarrow \Sigma(1385)\pi, F\text{-wave}$				$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.31 ± 0.04	ZHANG	13A	DPWA	Multichannel
−0.126 ± 0.055	<sup>3</sup> CAMERON	78	DPWA	$K^-p \rightarrow \Sigma(1385)\pi$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1890) \rightarrow N\bar{K}^*(892), S=1/2$				$(\Gamma_1\Gamma_7)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.17 ± 0.05	ZHANG	13A	DPWA	Multichannel
−0.07 ± 0.03	<sup>3,4</sup> CAMERON	78B	DPWA	$K^-p \rightarrow N\bar{K}^*$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1890) \rightarrow N\bar{K}^*(892), S=3/2, F\text{-wave}$				$(\Gamma_1\Gamma_8)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.11 ± 0.03	ZHANG	13A	DPWA	Multichannel

See key on page 601

# Baryon Particle Listings

## $\Lambda(1890)$ , $\Lambda(2000)$ , $\Lambda(2020)$

$(\Gamma_1 \Gamma_f)^\frac{1}{2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(1890) \rightarrow \Lambda\omega$	$(\Gamma_1 \Gamma_g)^\frac{1}{2} / \Gamma$
VALUE	DOCUMENT ID TECN COMMENT
seen	BACCARI 77 IPWA $K^- p \rightarrow \Lambda\omega$
0.032	<sup>2</sup> NAKKASYAN 75 DPWA $K^- p \rightarrow \Lambda\omega$

### $\Lambda(1890)$ FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> Found in one of two best solutions.  
<sup>3</sup> The published sign has been changed to be in accord with the baryon-first convention.  
<sup>4</sup> Upper limits on the  $P_3$  and  $F_3$  waves are each 0.03.

### $\Lambda(1890)$ REFERENCES

ZHANG 13A PR C88 035205	H. Zhang <i>et al.</i> (KSU)
PDG 82 PL 111B 1	M. Roos <i>et al.</i> (HELSE, CIT, CERN)
GOPAL 80 Toronto Conf. 159	G.P. Gopal (RHEL) IJP
ALSTON-... 78 PR D18 182	M. Alston-Garnjost <i>et al.</i> (LBL, MTHO+) IJP
Also PRL 38 1007	M. Alston-Garnjost <i>et al.</i> (LBL, MTHO+) IJP
CAMERON 78 NP B143 189	W. Cameron <i>et al.</i> (RHEL, LOIC) IJP
CAMERON 78B NP B146 327	B. Baccari <i>et al.</i> (RHEL, LOIC) IJP
BACCARI 77 NC 41A 96	G.P. Gopal <i>et al.</i> (SACL, CDEF) IJP
GOPAL 77 NP B119 362	B.R. Martin, M.K. Pidcock, R.G. Moorhouse (LOUC+) IJP
MARTIN 77 NP B127 349	B.R. Martin, M.K. Pidcock (LOUC) IJP
Also NP B126 266	B.R. Martin, M.K. Pidcock (LOUC) IJP
Also NP B126 285	R.J. Hemingway <i>et al.</i> (CERN, HEIDH, MPIM) IJP
HEMINGWAY 75 NP B91 12	A. Nakkasyan (CERN) IJP
NAKKASYAN 75 NP B93 85	

## $\Lambda(2000)$

$$I(J^P) = 0(?)^? \quad \text{Status: } *$$

OMITTED FROM SUMMARY TABLE

ZHANG 13A claims a  $J^P = 1/2^-$  state.

We list here all the ambiguous resonance possibilities with a mass around 2 GeV. The proposed quantum numbers are  $D_3$  (BARBARO-GALTIERI 70 in  $\Sigma\pi$ ),  $D_3+F_5$ ,  $P_3+D_5$ , or  $P_1+D_3$  (BRANDSTETER 72 in  $\Lambda\omega$ ), and  $S_1$  (CAMERON 78B in  $N\bar{K}^*$ ). The first two of the above analyses should now be considered obsolete. See also NAKKASYAN 75.

### $\Lambda(2000)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 2000</math> OUR ESTIMATE</b>			
2020 $\pm 16$	ZHANG 13A	DPWA	Multichannel
2030 $\pm 30$	CAMERON 78B	DPWA	$K^- p \rightarrow N\bar{K}^*$
1935 to 1971	<sup>1</sup> BRANDSTET...72	DPWA	$K^- p \rightarrow \Lambda\omega$
1951 to 2034	<sup>1</sup> BRANDSTET...72	DPWA	$K^- p \rightarrow \Lambda\omega$
2010 $\pm 30$	BARBARO-... 70	DPWA	$K^- p \rightarrow \Sigma\pi$

### $\Lambda(2000)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
255 $\pm 63$	ZHANG 13A	DPWA	Multichannel
125 $\pm 25$	CAMERON 78B	DPWA	$K^- p \rightarrow N\bar{K}^*$
180 to 240	<sup>1</sup> BRANDSTET...72	DPWA	(lower mass)
73 to 154	<sup>1</sup> BRANDSTET...72	DPWA	(higher mass)
130 $\pm 50$	BARBARO-... 70	DPWA	$K^- p \rightarrow \Sigma\pi$

### $\Lambda(2000)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	(27 $\pm$ 6) %
$\Gamma_2$ $\Sigma\pi$	
$\Gamma_3$ $\Lambda\eta$	(16 $\pm$ 7) %
$\Gamma_4$ $\Lambda\omega$	
$\Gamma_5$ $N\bar{K}^*(892)$ , $S=1/2$ , $S$ -wave	
$\Gamma_6$ $N\bar{K}^*(892)$ , $S=3/2$ , $D$ -wave	

### $\Lambda(2000)$ BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b>0.27 <math>\pm</math> 0.06</b>	ZHANG 13A DPWA Multichannel

$(\Gamma_1 \Gamma_f)^\frac{1}{2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2000) \rightarrow \Sigma\pi$	$(\Gamma_1 \Gamma_2)^\frac{1}{2} / \Gamma$
VALUE	DOCUMENT ID TECN COMMENT
-0.07 $\pm$ 0.03	ZHANG 13A DPWA Multichannel
-0.20 $\pm$ 0.04	BARBARO-... 70 DPWA $K^- p \rightarrow \Sigma\pi$

$\Gamma(\Lambda\eta)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b>0.16 <math>\pm</math> 0.07</b>	ZHANG 13A DPWA Multichannel

$(\Gamma_1 \Gamma_f)^\frac{1}{2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2000) \rightarrow \Lambda\omega$	$(\Gamma_1 \Gamma_g)^\frac{1}{2} / \Gamma$
VALUE	DOCUMENT ID TECN COMMENT
0.17 to 0.25	<sup>1</sup> BRANDSTET...72 DPWA (lower mass)
0.04 to 0.15	<sup>1</sup> BRANDSTET...72 DPWA (higher mass)

$(\Gamma_1 \Gamma_f)^\frac{1}{2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2000) \rightarrow N\bar{K}^*(892)$ , $S=1/2$ , $S$ -wave	$(\Gamma_1 \Gamma_g)^\frac{1}{2} / \Gamma$
VALUE	DOCUMENT ID TECN COMMENT
-0.12 $\pm$ 0.03	<sup>2</sup> CAMERON 78B DPWA $K^- p \rightarrow N\bar{K}^*$

$(\Gamma_1 \Gamma_f)^\frac{1}{2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2000) \rightarrow N\bar{K}^*(892)$ , $S=3/2$ , $D$ -wave	$(\Gamma_1 \Gamma_g)^\frac{1}{2} / \Gamma$
VALUE	DOCUMENT ID TECN COMMENT
+0.34 $\pm$ 0.05	ZHANG 13A DPWA Multichannel
+0.09 $\pm$ 0.03	CAMERON 78B DPWA $K^- p \rightarrow N\bar{K}^*$

### $\Lambda(2000)$ FOOTNOTES

- <sup>1</sup> The parameters quoted here are ranges from the three best fits; the lower state probably has  $J \leq 3/2$ , and the higher one probably has  $J \leq 5/2$ .  
<sup>2</sup> The published sign has been changed to be in accord with the baryon-first convention.

### $\Lambda(2000)$ REFERENCES

ZHANG 13A PR C88 035205	H. Zhang <i>et al.</i> (KSU)
CAMERON 78B NP B146 327	W. Cameron <i>et al.</i> (RHEL, LOIC) IJP
NAKKASYAN 75 NP B93 85	A. Nakkasyan (CERN) IJP
BRANDSTET... 72 NP B39 13	A.A. Brandstetter <i>et al.</i> (RHEL, CDEF+) IJP
BARBARO-... 70 Duke Conf. 173	A. Barbaro-Galtieri (LRL) IJP
Hyperon Resonances, 1970	

## $\Lambda(2020)$ 7/2<sup>+</sup>

$$I(J^P) = 0(\frac{7}{2}^+) \quad \text{Status: } *$$

OMITTED FROM SUMMARY TABLE

In LITCHFIELD 71, need for the state rests solely on a possibly inconsistent polarization measurement at 1.784 GeV/c. HEMINGWAY 75 does not require this state. GOPAL 77 does not need it in either  $N\bar{K}$  or  $\Sigma\pi$ . With new  $K^- n$  angular distributions included, DECLAIS 77 sees it. However, this and other new data are included in GOPAL 80 and the state is not required. BACCARI 77 weakly supports it.

### $\Lambda(2020)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 2020</math> OUR ESTIMATE</b>			
2043 $\pm 22$	ZHANG 13A	DPWA	Multichannel
2140	BACCARI 77	DPWA	$K^- p \rightarrow \Lambda\omega$
2117	DECLAIS 77	DPWA	$\bar{K}N \rightarrow \bar{K}N$
2100 $\pm 30$	LITCHFIELD 71	DPWA	$K^- p \rightarrow \bar{K}N$
2020 $\pm 20$	BARBARO-... 70	DPWA	$K^- p \rightarrow \Sigma\pi$

### $\Lambda(2020)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
200 $\pm 75$	ZHANG 13A	DPWA	Multichannel
128	BACCARI 77	DPWA	$K^- p \rightarrow \Lambda\omega$
167	DECLAIS 77	DPWA	$\bar{K}N \rightarrow \bar{K}N$
120 $\pm 30$	LITCHFIELD 71	DPWA	$K^- p \rightarrow \bar{K}N$
160 $\pm 30$	BARBARO-... 70	DPWA	$K^- p \rightarrow \Sigma\pi$

### $\Lambda(2020)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	
$\Gamma_2$ $\Sigma\pi$	
$\Gamma_3$ $\Lambda\omega$	
$\Gamma_4$ $N\bar{K}^*(892)$ , $S=1/2$	(30 $\pm$ 9) %

### $\Lambda(2020)$ BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
0.028 $\pm$ 0.005	ZHANG 13A DPWA Multichannel
0.05	DECLAIS 77 DPWA $\bar{K}N \rightarrow \bar{K}N$
0.05 $\pm$ 0.02	LITCHFIELD 71 DPWA $K^- p \rightarrow \bar{K}N$

# Baryon Particle Listings

## $\Lambda(2020)$ , $\Lambda(2050)$ , $\Lambda(2100)$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2020) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
+0.02±0.01	ZHANG	13A	DPWA Multichannel
−0.15±0.02	BARBARO...	70	DPWA $K^-p \rightarrow \Sigma\pi$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2020) \rightarrow \Lambda\omega$	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<0.05	BACCARI	77	DPWA $K^-p \rightarrow \Lambda\omega$

$\Gamma(N\bar{K}^*(892), S=1/2)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.30±0.09</b>	ZHA NG	13A	DPWA Multichannel

### $\Lambda(2020)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL)
BACCARI	77	NC 41A 96	B. Baccari <i>et al.</i>	(SACL, CDEF) IJP
DECLAIS	77	CERN 77-16	Y. Declais <i>et al.</i>	(CAEN, CERN) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL)
HEMINGWAY	75	NP B91 12	R.J. Hemingway <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
LITCHFIELD	71	NP B30 125	P.J. Litchfield <i>et al.</i>	(RHEL, CDEF, SACL) IJP
BARBARO...	70	Duke Conf. 173	A. Barbaro-Galtieri	(LRL) IJP
Hyperon Resonances, 1970				

$\Lambda(2050)$ 3/2 <sup>−</sup>	$I(J^P) = 0(\frac{3}{2}^-)$ Status: *
OMITTED FROM SUMMARY TABLE	

### $\Lambda(2050)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2056±22</b>	ZHA NG	13A	DPWA Multichannel

### $\Lambda(2050)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>493±61</b>	ZHA NG	13A	DPWA Multichannel

### $\Lambda(2050)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	(19 ±4 ) %
$\Gamma_2$ $\Sigma\pi$	( 6.0±3.0) %
$\Gamma_3$ $\Sigma^*(1385)\pi$ , S-wave	( 8 ±6 ) %
$\Gamma_4$ $\Sigma^*(1385)\pi$ , D-wave	( 4.0±3.0) %
$\Gamma_5$ $N\bar{K}^*(892)$ , S=1/2	(23 ±7 ) %

### $\Lambda(2050)$ BRANCHING RATIOS

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.19±0.04</b>	ZHA NG	13A	DPWA Multichannel

$\Gamma(\Sigma\pi)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.06±0.03</b>	ZHA NG	13A	DPWA Multichannel

$\Gamma(\Sigma^*(1385)\pi, S\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.08±0.06</b>	ZHA NG	13A	DPWA Multichannel

$\Gamma(\Sigma^*(1385)\pi, D\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.04±0.03</b>	ZHA NG	13A	DPWA Multichannel

$\Gamma(N\bar{K}^*(892), S=1/2)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.23±0.07</b>	ZHA NG	13A	DPWA Multichannel

### $\Lambda(2050)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
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$\Lambda(2100)$ 7/2 <sup>−</sup>	$I(J^P) = 0(\frac{7}{2}^-)$ Status: ****
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Most of the results published before 1973 are now obsolete and have been omitted. They may be found in our 1982 edition Physics Letters **111B** 1 (1982).

This entry only includes results from partial-wave analyses. Parameters of peaks seen in cross sections and in invariant-mass distributions around 2100 MeV used to be listed in a separate entry immediately following. It may be found in our 1986 edition Physics Letters **170B** 1 (1986).

### $\Lambda(2100)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2090 to 2110 (≈ 2100) OUR ESTIMATE</b>			
2086± 6	ZHANG	13A	DPWA Multichannel
2104±10	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
2106±30	DEBELLEFON	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
2110±10	GOPAL	77	DPWA $\bar{K}N$ multichannel
2105±10	HEMINGWAY	75	DPWA $K^-p \rightarrow \bar{K}N$
2115±10	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2094	BACCARI	77	DPWA $K^-p \rightarrow \Lambda\omega$
2094	DECLAIS	77	DPWA $\bar{K}N \rightarrow \bar{K}N$
2110 or 2089	<sup>1</sup> NAKKASYAN	75	DPWA $K^-p \rightarrow \Lambda\omega$

### $\Lambda(2100)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>100 to 250 (≈ 200) OUR ESTIMATE</b>			
305±16	ZHANG	13A	DPWA Multichannel
157±40	DEBELLEFON	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
250±30	GOPAL	77	DPWA $\bar{K}N$ multichannel
241±30	HEMINGWAY	75	DPWA $K^-p \rightarrow \bar{K}N$
152±15	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
98	BACCARI	77	DPWA $K^-p \rightarrow \Lambda\omega$
250	DECLAIS	77	DPWA $\bar{K}N \rightarrow \bar{K}N$
244 or 302	<sup>1</sup> NAKKASYAN	75	DPWA $K^-p \rightarrow \Lambda\omega$

### $\Lambda(2100)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2023	ZHANG	13A	DPWA Multichannel

#### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
239	ZHANG	13A	DPWA Multichannel

### $\Lambda(2100)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	25–35 %
$\Gamma_2$ $\Sigma\pi$	~ 5 %
$\Gamma_3$ $\Lambda\eta$	<3 %
$\Gamma_4$ $\Xi K$	<3 %
$\Gamma_5$ $\Lambda\omega$	<8 %
$\Gamma_6$ $N\bar{K}^*(892)$	10–20 %
$\Gamma_7$ $N\bar{K}^*(892)$ , S=3/2, D-wave	
$\Gamma_8$ $N\bar{K}^*(892)$ , S=1/2, G-wave	
$\Gamma_9$ $N\bar{K}^*(892)$ , S=3/2, G-wave	

The above branching fractions are our estimates, not fits or averages.

### $\Lambda(2100)$ BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.25 to 0.35 OUR ESTIMATE</b>			
0.23±0.01	ZHANG	13A	DPWA Multichannel
0.34±0.03	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
0.24±0.06	DEBELLEFON	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
0.31±0.03	HEMINGWAY	75	DPWA $K^-p \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.29	DECLAIS	77	DPWA $\bar{K}N \rightarrow \bar{K}N$
0.30±0.03	GOPAL	77	DPWA See GOPAL 80

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
+0.03±0.01	ZHANG	13A	DPWA Multichannel
+0.12±0.04	GOPAL	77	DPWA $\bar{K}N$ multichannel
+0.11±0.01	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$

See key on page 601

# Baryon Particle Listings

## $\Lambda(2100)$ , $\Lambda(2110)$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow \Lambda\eta$	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$-0.050 \pm 0.020$	RADER 73 MPWA $K^-p \rightarrow \Lambda\eta$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow \Xi K$	$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$0.035 \pm 0.018$	LITCHFIELD 71 DPWA $K^-p \rightarrow \Xi K$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.003	MULLER 69B DPWA $K^-p \rightarrow \Xi K$
0.05	TRIPP 67 RVUE $K^-p \rightarrow \Xi K$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow \Lambda\omega$	$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$-0.070$	<sup>2</sup> BACCARI 77 DPWA $GD_{37}$ wave
$+0.011$	<sup>2</sup> BACCARI 77 DPWA $GG_{17}$ wave
$+0.008$	<sup>2</sup> BACCARI 77 DPWA $GG_{37}$ wave
0.122 or 0.154	<sup>1</sup> NAKKASYAN 75 DPWA $K^-p \rightarrow \Lambda\omega$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow N\bar{K}^*(892), S=3/2, D\text{-wave}$	$(\Gamma_1\Gamma_7)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$+0.16 \pm 0.02$	ZHA NG 13A DPWA Multichannel
$+0.21 \pm 0.04$	CAMERON 78B DPWA $K^-p \rightarrow N\bar{K}^*$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow N\bar{K}^*(892), S=1/2, G\text{-wave}$	$(\Gamma_1\Gamma_8)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$-0.03 \pm 0.02$	ZHANG 13A DPWA Multichannel
$-0.04 \pm 0.03$	<sup>3</sup> CAMERON 78B DPWA $K^-p \rightarrow N\bar{K}^*$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2100) \rightarrow N\bar{K}^*(892), S=3/2, G\text{-wave}$	$(\Gamma_1\Gamma_9)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$+0.08 \pm 0.02$	ZHA NG 13A DPWA Multichannel

### $\Lambda(2100)$ FOOTNOTES

- <sup>1</sup> The NAKKASYAN 75 values are from the two best solutions found. Each has the  $\Lambda(2100)$  and one additional resonance ( $P_3$  or  $F_5$ ).  
<sup>2</sup> Note that the three for BACCARI 77 entries are for three different waves.  
<sup>3</sup> The published sign has been changed to be in accord with the baryon-first convention. The upper limit on the  $G_3$  wave is 0.03.

### $\Lambda(2100)$ REFERENCES

ZHANG 13A PR C08 035205	H. Zhang <i>et al.</i> (KSU)
PDG 86 PL 170B 1	M. Aguilar-Benitez <i>et al.</i> (CERN, CIT+)
PDG 82 PL 111B 1	M. Roos <i>et al.</i> (HELSE, CIT, CERN)
GOPAL 80 Toronto Conf. 159	G.P. Gopal (RHEL) IJP
CAMERON 78B NP B146 327	W. Cameron <i>et al.</i> (RHEL, LOIC) IJP
DEBELLEFON 78 NC 42A 403	A. de Bellefon <i>et al.</i> (CDEF, SACL) IJP
BACCARI 77 NC 41A 96	B. Baccari <i>et al.</i> (SACL, CDEF) IJP
DECLAIS 77 CERN 77-16	Y. Declais <i>et al.</i> (CAEN, CERN) IJP
GOPAL 77 NP B119 362	G.P. Gopal <i>et al.</i> (LOIC, RHEL) IJP
HEMINGWAY 75 NP B91 12	R.J. Hemingway <i>et al.</i> (CERN, HEIDH, MPIM) IJP
NAKKASYAN 75 NP B93 85	A. Nakkasyan (CERN) IJP
KANE 74 LBL-2452	D.F. Kane (LBL) IJP
RADER 73 NC 16A 178	R.K. Rader <i>et al.</i> (SACL, HEID, CERN+) IJP
LITCHFIELD 71 NP B30 125	P.J. Litchfield <i>et al.</i> (RHEL, CDEF, SACL) IJP
MULLER 69B Thesis UCRL 19372	R.A. Muller (LRL)
TRIPP 67 NP B3 10	R.D. Tripp <i>et al.</i> (LRL, SLAC, CERN+)

$\Lambda(2110) \ 5/2^+$

$I(J^P) = 0(\frac{5}{2}^+)$  Status: \*\*\*

For results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982). All the references have been retained.

This resonance is in the Baryon Summary Table, but the evidence for it could be better.

### $\Lambda(2110)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2090 to 2140 (<math>\approx 2110</math>) OUR ESTIMATE</b>			
$2036 \pm 13$	ZHA NG 13A DPWA Multichannel		
$2092 \pm 25$	GOPAL 80 DPWA $\bar{K}N \rightarrow \bar{K}N$		
$2125 \pm 25$	CAMERON 78B DPWA $K^-p \rightarrow N\bar{K}^*$		
$2106 \pm 50$	DEBELLEFON 78 DPWA $\bar{K}N \rightarrow \bar{K}N$		
$2140 \pm 20$	DEBELLEFON 77 DPWA $K^-p \rightarrow \Sigma\pi$		
$2100 \pm 50$	GOPAL 77 DPWA $\bar{K}N$ multichannel		
$2112 \pm 7$	KANE 74 DPWA $K^-p \rightarrow \Sigma\pi$		
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2137	BACCARI 77 DPWA $K^-p \rightarrow \Lambda\omega$		
2103	<sup>1</sup> NAKKASYAN 75 DPWA $K^-p \rightarrow \Lambda\omega$		

### $\Lambda(2110)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>150 to 250 (<math>\approx 200</math>) OUR ESTIMATE</b>			
$400 \pm 38$	ZHANG 13A DPWA Multichannel		
$245 \pm 25$	GOPAL 80 DPWA $\bar{K}N \rightarrow \bar{K}N$		
$160 \pm 30$	CAMERON 78B DPWA $K^-p \rightarrow N\bar{K}^*$		
$251 \pm 50$	DEBELLEFON 78 DPWA $\bar{K}N \rightarrow \bar{K}N$		
$140 \pm 20$	DEBELLEFON 77 DPWA $K^-p \rightarrow \Sigma\pi$		
$200 \pm 50$	GOPAL 77 DPWA $\bar{K}N$ multichannel		
$190 \pm 30$	KANE 74 DPWA $K^-p \rightarrow \Sigma\pi$		
• • • We do not use the following data for averages, fits, limits, etc. • • •			
132	BACCARI 77 DPWA $K^-p \rightarrow \Lambda\omega$		
391	<sup>1</sup> NAKKASYAN 75 DPWA $K^-p \rightarrow \Lambda\omega$		

### $\Lambda(2110)$ POLE POSITION

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1970	ZHANG 13A DPWA Multichannel		

### $-2\times$ IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
350	ZHANG 13A DPWA Multichannel		

### $\Lambda(2110)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \ N\bar{K}$	5–25 %
$\Gamma_2 \ \Sigma\pi$	10–40 %
$\Gamma_3 \ \Lambda\omega$	seen
$\Gamma_4 \ \Sigma(1385)\pi$	seen
$\Gamma_5 \ \Sigma(1385)\pi, P\text{-wave}$	
$\Gamma_6 \ N\bar{K}^*(892)$	10–60 %
$\Gamma_7 \ N\bar{K}^*(892), S=1/2$	
$\Gamma_8 \ N\bar{K}^*(892), S=3/2, P\text{-wave}$	

The above branching fractions are our estimates, not fits or averages.

### $\Lambda(2110)$ BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID
<b>0.05 to 0.25 OUR ESTIMATE</b>	
$0.083 \pm 0.005$	ZHANG 13A DPWA Multichannel
$0.07 \pm 0.03$	GOPAL 80 DPWA $\bar{K}N \rightarrow \bar{K}N$
$0.27 \pm 0.06$	<sup>2</sup> DEBELLEFON 78 DPWA $\bar{K}N \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.07 \pm 0.03$	GOPAL 77 DPWA See GOPAL 80

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2110) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$+0.04 \pm 0.01$	ZHANG 13A DPWA Multichannel
$+0.14 \pm 0.01$	DEBELLEFON 77 DPWA $K^-p \rightarrow \Sigma\pi$
$+0.20 \pm 0.03$	KANE 74 DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$+0.10 \pm 0.03$	GOPAL 77 DPWA $\bar{K}N$ multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2110) \rightarrow \Lambda\omega$	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$<0.05$	BACCARI 77 DPWA $K^-p \rightarrow \Lambda\omega$
0.112	<sup>1</sup> NAKKASYAN 75 DPWA $K^-p \rightarrow \Lambda\omega$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2110) \rightarrow \Sigma(1385)\pi, P\text{-wave}$	$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$+0.04 \pm 0.01$	ZHANG 13A DPWA Multichannel
$+0.071 \pm 0.025$	<sup>3</sup> CAMERON 78 DPWA $K^-p \rightarrow \Sigma(1385)\pi$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2110) \rightarrow N\bar{K}^*(892), S=1/2$	$(\Gamma_1\Gamma_7)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$-0.09 \pm 0.01$	ZHANG 13A DPWA Multichannel
$-0.17 \pm 0.04$	<sup>4</sup> CAMERON 78B DPWA $K^-p \rightarrow N\bar{K}^*$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2110) \rightarrow N\bar{K}^*(892), S=3/2, P\text{-wave}$	$(\Gamma_1\Gamma_8)^{1/2}/\Gamma$
VALUE	DOCUMENT ID
$0.24 \pm 0.01$	ZHANG 13A DPWA Multichannel

# Baryon Particle Listings

## $\Lambda(2110)$ , $\Lambda(2325)$ , $\Lambda(2350)$ , $\Lambda(2585)$ Bumps

### $\Lambda(2110)$ FOOTNOTES

- <sup>1</sup> Found in one of two best solutions.  
<sup>2</sup> The published error of 0.6 was a misprint.  
<sup>3</sup> The CAMERON 78 upper limit on  $F$ -wave decay is 0.03. The sign here has been changed to be in accord with the baryon-first convention.  
<sup>4</sup> The published sign has been changed to be in accord with the baryon-first convention. The CAMERON 78B upper limits on the  $P_3$  and  $F_3$  waves are each 0.03.

### $\Lambda(2110)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HEL5, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
CAMERON	78	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
CAMERON	78B	NP B146 327	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
DEBELLEFON	78	NC 42A 403	A. de Bellefon <i>et al.</i>	(CDEF, SACL) IJP
BACCARI	77	NC 41A 96	B. Baccari <i>et al.</i>	(SACL, CDEF) IJP
DEBELLEFON	77	NC 37A 175	A. de Bellefon <i>et al.</i>	(CDEF, SACL) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
NAKKASYAN	75	NP B93 85	A. Nakkasyan	(CERN) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP

## $\Lambda(2325)$ $3/2^-$

$$I(J^P) = 0(\frac{3}{2}^-) \text{ Status: } *$$

OMITTED FROM SUMMARY TABLE

BACCARI 77 finds this state with either  $J^P = 3/2^-$  or  $3/2^+$  in a energy-dependent partial-wave analyses of  $K^-p \rightarrow \Lambda\omega$  from 2070 to 2436 MeV. A subsequent semi-energy-independent analysis from threshold to 2436 MeV selects  $3/2^-$ . DEBELLEFON 78 (same group) also sees this state in an energy-dependent partial-wave analysis of  $K^-p \rightarrow \bar{K}N$  data, and finds  $J^P = 3/2^-$  or  $3/2^+$ . They again prefer  $J^P = 3/2^-$ , but only on the basis of model-dependent considerations.

### $\Lambda(2325)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 2325$ OUR ESTIMATE			
2342 $\pm$ 30	DEBELLEFON 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
2327 $\pm$ 20	BACCARI 77	DPWA	$K^-p \rightarrow \Lambda\omega$

### $\Lambda(2325)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
177 $\pm$ 40	DEBELLEFON 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
160 $\pm$ 40	BACCARI 77	IPWA	$K^-p \rightarrow \Lambda\omega$

### $\Lambda(2325)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	$\sim 12\%$
$\Gamma_2$ $\Lambda\omega$	$\sim 10\%$

### $\Lambda(2325)$ BRANCHING RATIOS

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
0.19 $\pm$ 0.06	DEBELLEFON 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$	
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2325) \rightarrow \Lambda\omega$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
0.06 $\pm$ 0.02	<sup>1</sup> BACCARI 77	IPWA	$DS_{33}$ wave	
0.05 $\pm$ 0.02	<sup>1</sup> BACCARI 77	DPWA	$DD_{13}$ wave	
0.08 $\pm$ 0.03	<sup>1</sup> BACCARI 77	DPWA	$DD_{33}$ wave	

### $\Lambda(2325)$ FOOTNOTES

- <sup>1</sup> Note that the three BACCARI 77 entries are for three different waves.

### $\Lambda(2325)$ REFERENCES

DEBELLEFON 78	NC 42A 403	A. de Bellefon <i>et al.</i>	(CDEF, SACL) IJP
BACCARI 77	NC 41A 96	B. Baccari <i>et al.</i>	(SACL, CDEF) IJP

## $\Lambda(2350)$ $9/2^+$

$$I(J^P) = 0(\frac{9}{2}^+) \text{ Status: } ***$$

DAUM 68 favors  $J^P = 7/2^-$  or  $9/2^+$ . BRICMAN 70 favors  $9/2^+$ . LASINSKI 71 suggests three states in this region using a Pomeron + resonances model. There are now also three formation experiments from the College de France-Saclay group, DEBELLEFON 77, BACCARI 77, and DEBELLEFON 78, which find  $9/2^+$  in energy-dependent partial-wave analyses of  $\bar{K}N \rightarrow \Sigma\pi$ ,  $\Lambda\omega$ , and  $N\bar{K}$ .

### $\Lambda(2350)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2340 to 2370 (<math>\approx 2350</math>) OUR ESTIMATE</b>			
2370 $\pm$ 50	DEBELLEFON 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
2365 $\pm$ 20	DEBELLEFON 77	DPWA	$K^-p \rightarrow \Sigma\pi$
2358 $\pm$ 6	BRICMAN 70	CNTR	Total, charge exchange
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2372	BACCARI 77	DPWA	$K^-p \rightarrow \Lambda\omega$
2344 $\pm$ 15	COOL 70	CNTR	$K^-p$ , $K^-d$ total
2360 $\pm$ 20	LU 70	CNTR	$\gamma p \rightarrow K^+Y^*$
2340 $\pm$ 7	BUGG 68	CNTR	$K^-p$ , $K^-d$ total

### $\Lambda(2350)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>100 to 250 (<math>\approx 150</math>) OUR ESTIMATE</b>			
204 $\pm$ 50	DEBELLEFON 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
110 $\pm$ 20	DEBELLEFON 77	DPWA	$K^-p \rightarrow \Sigma\pi$
324 $\pm$ 30	BRICMAN 70	CNTR	Total, charge exchange
• • • We do not use the following data for averages, fits, limits, etc. • • •			
257	BACCARI 77	DPWA	$K^-p \rightarrow \Lambda\omega$
190	COOL 70	CNTR	$K^-p$ , $K^-d$ total
55	LU 70	CNTR	$\gamma p \rightarrow K^+Y^*$
140 $\pm$ 20	BUGG 68	CNTR	$K^-p$ , $K^-d$ total

### $\Lambda(2350)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\bar{K}$	$\sim 12\%$
$\Gamma_2$ $\Sigma\pi$	$\sim 10\%$
$\Gamma_3$ $\Lambda\omega$	

The above branching fractions are our estimates, not fits or averages.

### $\Lambda(2350)$ BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>VALUE</b>				
$\sim 0.12$ OUR ESTIMATE				
0.12 $\pm$ 0.04	DEBELLEFON 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$	
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2350) \rightarrow \Sigma\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
<b>VALUE</b>				
$-0.11\pm 0.02$	DEBELLEFON 77	DPWA	$K^-p \rightarrow \Sigma\pi$	
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Lambda(2350) \rightarrow \Lambda\omega$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
<b>VALUE</b>				
$<0.05$	BACCARI 77	DPWA	$K^-p \rightarrow \Lambda\omega$	

### $\Lambda(2350)$ REFERENCES

DEBELLEFON 78	NC 42A 403	A. de Bellefon <i>et al.</i>	(CDEF, SACL) IJP
BACCARI 77	NC 41A 96	B. Baccari <i>et al.</i>	(SACL, CDEF) IJP
DEBELLEFON 77	NC 37A 175	A. de Bellefon <i>et al.</i>	(CDEF, SACL) IJP
LASINSKI 71	NP B29 125	T.A. Lasinski	(EF) IJP
BRICMAN 70	PL 31B 152	C. Bricman <i>et al.</i>	(CERN, CAEN, SACL)
COOL 70	PR D1 1887	R.L. Cool <i>et al.</i>	(BNL) I
Also	PR L16 1228	R.L. Cool <i>et al.</i>	(BNL) I
LU 70	PR D2 1846	D.C. Lu <i>et al.</i>	(YALE)
BUGG 68	PR 168 1466	D.V. Bugg <i>et al.</i>	(RHEL, BIRM, CAVE) I
DAUM 68	NP B7 19	C. Daum <i>et al.</i>	(CERN) JP

## $\Lambda(2585)$ Bumps

$$I(J^P) = 0(?)^? \text{ Status: } **$$

OMITTED FROM SUMMARY TABLE

### $\Lambda(2585)$ MASS (BUMPS)

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 2585</math> OUR ESTIMATE</b>			
2585 $\pm$ 45	ABRAMS 70	CNTR	$K^-p$ , $K^-d$ total
2530 $\pm$ 25	LU 70	CNTR	$\gamma p \rightarrow K^+Y^*$

### $\Lambda(2585)$ WIDTH (BUMPS)

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
300	ABRAMS 70	CNTR	$K^-p$ , $K^-d$ total
150	LU 70	CNTR	$\gamma p \rightarrow K^+Y^*$



$\Lambda(2585)$ DECAY MODES (BUMPS)			
Mode			
$\Gamma_1$	$N\overline{K}$		

$\Lambda(2585)$ BRANCHING RATIOS (BUMPS)			
$(J+\frac{1}{2})\times\Gamma(N\overline{K})/\Gamma_{\text{total}}$			$\Gamma_1/\Gamma$
$J$ is not known, so only $(J+\frac{1}{2})\times\Gamma(N\overline{K})/\Gamma_{\text{total}}$ can be given.			
VALUE	DOCUMENT ID	TECN	COMMENT
1	ABRAMS	70	CNTR $K^-p, K^-d$ total
$0.12\pm0.12$	<sup>1</sup> BRICMAN	70	CNTR Total, charge exchange

$\Lambda(2585)$ FOOTNOTES (BUMPS)			
<sup>1</sup> The resonance is at the end of the region analyzed — no clear signal.			

$\Lambda(2585)$ REFERENCES (BUMPS)			
ABRAMS	70	PR D1 1917	R.J. Abrams <i>et al.</i> (BNL) I
Also		PRL 16 1228	R.L. Cool <i>et al.</i> (BNL) I
BRICMAN	70	PL 31B 152	C. Bricman <i>et al.</i> (CERN, CAEN, SACL)
LU	70	PR D2 1846	D.C. Lu <i>et al.</i> (YALE)

Baryon Particle Listings

$\Sigma^+$

$\Sigma$  BARYONS  
( $S = -1, I = 1$ )  
 $\Sigma^+ = uus, \Sigma^0 = uds, \Sigma^- = dds$

$\Sigma^+$

$I(J^P) = 1(\frac{1}{2}^+)$  Status: \*\*\*

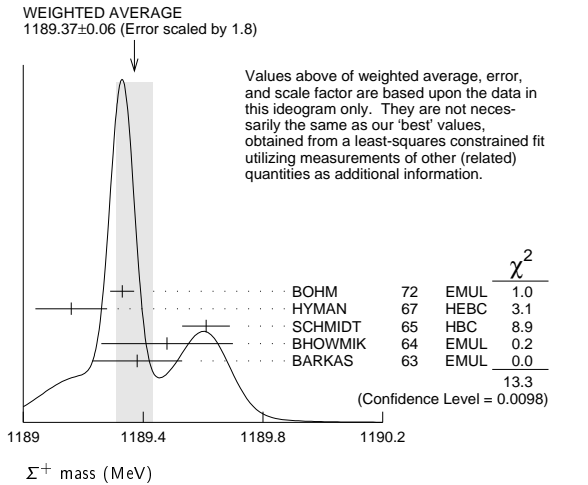
We have omitted some results that have been superseded by later experiments. See our earlier editions.

$\Sigma^+$  MASS

The fit uses  $\Sigma^+, \Sigma^0, \Sigma^-$ , and  $\Lambda$  mass and mass-difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1189.37 ± 0.07 OUR FIT</b>				Error includes scale factor of 2.2.
<b>1189.37 ± 0.06 OUR AVERAGE</b>				Error includes scale factor of 1.8. See the ideogram below.
1189.33 ± 0.04	607	<sup>1</sup> BOHM	72	EMUL
1189.16 ± 0.12		HYMAN	67	HEBC
1189.61 ± 0.08	4205	SCHMIDT	65	HBC See note with $\Lambda$ mass
1189.48 ± 0.22	58	<sup>2</sup> BHOWMIK	64	EMUL
1189.38 ± 0.15	144	<sup>2</sup> BARKAS	63	EMUL

<sup>1</sup> BOHM 72 is updated with our 1973  $K^-, \pi^-,$  and  $\pi^0$  masses (Reviews of Modern Physics **45** S1 (1973)).  
<sup>2</sup> These masses have been raised 30 keV to take into account a 46 keV increase in the proton mass and a 21 keV decrease in the  $\pi^0$  mass (note added 1967 edition, Reviews of Modern Physics **39** 1 (1967)).



$\Sigma^+$  MEAN LIFE

Measurements with fewer than 1000 events have been omitted.

VALUE ( $10^{-10}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.8018 ± 0.0026 OUR AVERAGE</b>				
0.8038 ± 0.0040 ± 0.0014		BARBOSA	00	E761 hyperons, 375 GeV
0.8043 ± 0.0080 ± 0.0014		<sup>3</sup> BARBOSA	00	E761 hyperons, 375 GeV
0.798 ± 0.005	30k	MARRAFFINO	80	HBC $K^- p$ 0.42–0.5 GeV/c
0.807 ± 0.013	5719	CONFORTO	76	HBC $K^- p$ 1–1.4 GeV/c
0.795 ± 0.010	20k	EISELE	70	HBC $K^- p$ at rest
0.803 ± 0.008	10664	BARLOUTAUD	69	HBC $K^- p$ 0.4–1.2 GeV/c
0.83 ± 0.032	1300	<sup>4</sup> CHANG	66	HBC

<sup>3</sup> This is a measurement of the  $\Sigma^-$  lifetime. Here we assume  $CPT$  invariance; see below for the fractional  $\Sigma^+ - \Sigma^-$  lifetime difference obtained by BARBOSA 00.  
<sup>4</sup> We have increased the CHANG 66 error of 0.018; see our 1970 edition, Reviews of Modern Physics **42** 87 (1970).

$(\tau_{\Sigma^+} - \tau_{\Sigma^-}) / \tau_{\Sigma^+}$

A test of  $CPT$  invariance.

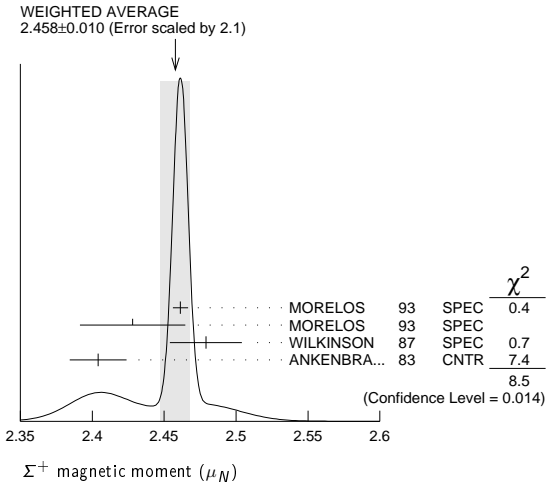
VALUE	DOCUMENT ID	TECN	COMMENT
<b>(-6 ± 12) × 10<sup>-4</sup></b>	BARBOSA	00	E761 hyperons, 375 GeV

$\Sigma^+$  MAGNETIC MOMENT

See the "Note on Baryon Magnetic Moments" in the  $\Lambda$  Listings. Measurements with an error  $\geq 0.1 \mu_N$  have been omitted.

VALUE ( $\mu_N$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.458 ± 0.010 OUR AVERAGE</b>				Error includes scale factor of 2.1. See the ideogram below.
2.4613 ± 0.0034 ± 0.0040	250k	MORELOS	93	SPEC $p$ Cu 800 GeV
2.428 ± 0.036 ± 0.007	12k	<sup>5</sup> MORELOS	93	SPEC $p$ Cu 800 GeV
2.479 ± 0.012 ± 0.022	137k	WILKINSON	87	SPEC $p$ Be 400 GeV
2.4040 ± 0.0198	44k	<sup>6</sup> ANKENBRA...	83	CNTR $p$ Cu 400 GeV

<sup>5</sup> We assume  $CPT$  invariance: this is (minus) the  $\Sigma^-$  magnetic moment as measured by MORELOS 93. See below for the moment difference testing  $CPT$ .  
<sup>6</sup> ANKENBRANDT 83 gives the value  $2.38 \pm 0.02 \mu_N$ . MORELOS 93 uses the same hyperon magnet and channel and claims to determine the field integral better, leading to the revised value given here.



$(\mu_{\Sigma^+} + \mu_{\Sigma^-}) / \mu_{\Sigma^+}$

A test of  $CPT$  invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.014 ± 0.015</b>	<sup>7</sup> MORELOS	93	SPEC $p$ Cu 800 GeV

<sup>7</sup> This is our calculation from the MORELOS 93 measurements of the  $\Sigma^+$  and  $\Sigma^-$  magnetic moments given above. The statistical error on  $\mu_{\Sigma^-}$  dominates the error here.

$\Sigma^+$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $p \pi^0$	(51.57 ± 0.30) %	
$\Gamma_2$ $n \pi^+$	(48.31 ± 0.30) %	
$\Gamma_3$ $p \gamma$	( 1.23 ± 0.05 ) × 10 <sup>-3</sup>	
$\Gamma_4$ $n \pi^+ \gamma$	[a] ( 4.5 ± 0.5 ) × 10 <sup>-4</sup>	
$\Gamma_5$ $\Lambda e^+ \nu_e$	( 2.0 ± 0.5 ) × 10 <sup>-5</sup>	

$\Delta S = \Delta Q$  (SQ) violating modes or  
 $\Delta S = 1$  weak neutral current (S1) modes

$\Gamma_6$ $n e^+ \nu_e$	SQ	< 5	× 10 <sup>-6</sup>	90%
$\Gamma_7$ $n \mu^+ \nu_\mu$	SQ	< 3.0	× 10 <sup>-5</sup>	90%
$\Gamma_8$ $p e^+ e^-$	S1	< 7	× 10 <sup>-6</sup>	
$\Gamma_9$ $p \mu^+ \mu^-$	S1	( 9 <sup>+9</sup> / <sub>-8</sub> )	× 10 <sup>-8</sup>	

[a] See the Listings below for the pion momentum range used in this measurement.

CONSTRAINED FIT INFORMATION

An overall fit to 2 branching ratios uses 14 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 7.7$  for 12 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-100	
$x_3$	12	-14
	$x_1$	$x_2$

See key on page 601

## Baryon Particle Listings

 $\Sigma^+$  $\Sigma^+$  BRANCHING RATIOS

$\Gamma(n\pi^+)/\Gamma(N\pi)$		$\Gamma_2/(\Gamma_1+\Gamma_2)$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.4836±0.0030 OUR FIT</b>			
<b>0.4836±0.0030 OUR AVERAGE</b>			
0.4828±0.0036	10k	<sup>8</sup> MARRAFFINO 80	HBC $K^- p$ 0.42–0.5 GeV/c
0.488 ±0.008	1861	NOWAK 78	HBC
0.484 ±0.015	537	TOVEE 71	EMUL
0.488 ±0.010	1331	BARLOUTAUD 69	HBC $K^- p$ 0.4–1.2 GeV/c
0.46 ±0.02	534	CHANG 66	HBC
0.490 ±0.024	308	HUMPHREY 62	HBC

<sup>8</sup> MARRAFFINO 80 actually gives  $\Gamma(p\pi^0)/\Gamma(\text{total}) = 0.5172 \pm 0.0036$ .

$\Gamma(p\gamma)/\Gamma(p\pi^0)$		$\Gamma_3/\Gamma_1$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>2.38±0.10 OUR FIT</b>			
<b>2.38±0.10 OUR AVERAGE</b>			
2.32±0.11±0.10	32k	TIMM 95	E761 $\Sigma^+$ 375 GeV
2.81±0.39+ <sup>0.21</sup> <sub>−0.43</sub>	408	HESSEY 89	CNTR $K^- p \rightarrow \Sigma^+ \pi^-$ at rest
2.52±0.28	190	<sup>9</sup> KOBAYASHI 87	CNTR $\pi^+ p \rightarrow \Sigma^+ K^+$
2.46+ <sup>0.30</sup> <sub>−0.35</sub>	155	BIAGI 85	CNTR CERN hyperon beam
2.11±0.38	46	MANZ 80	HBC $K^- p \rightarrow \Sigma^+ \pi^-$
2.1 ±0.3	45	ANG 69B	HBC $K^- p$ at rest
2.76±0.51	31	GERSHWIN 69B	HBC $K^- p \rightarrow \Sigma^+ \pi^-$
3.7 ±0.8	24	BAZIN 65	HBC $K^- p$ at rest

<sup>9</sup> KOBAYASHI 87 actually gives  $\Gamma(p\gamma)/\Gamma(\text{total}) = (1.30 \pm 0.15) \times 10^{-3}$ .

$\Gamma(n\pi^+\gamma)/\Gamma(n\pi^+)$		$\Gamma_4/\Gamma_2$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.93±0.10</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.27±0.05	29	ANG 69B	HBC $\pi^+ < 110$ MeV/c
~1.8		BAZIN 65B	HBC $\pi^+ < 116$ MeV/c

$\Gamma(\Lambda e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_5/\Gamma$	
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN COMMENT
<b>2.0±0.5 OUR AVERAGE</b>			
1.6±0.7	5	BALTAY 69	HBC $K^- p$ at rest
2.9±1.0	10	EISELE 69	HBC $K^- p$ at rest
2.0±0.8	6	BARASH 67	HBC $K^- p$ at rest

$\Gamma(ne^+\nu_e)/\Gamma(n\pi^+)$				$\Gamma_6/\Gamma_2$	
Test of $\Delta S = \Delta Q$ rule. Experiments with an effective denominator less than 100,000 have been omitted.					
<u>EFFECTIVE DENOM.</u> <u>EVTS</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt; 1.1 × 10<sup>-5</sup> OUR LIMIT</b> Our 90% CL limit = (2.3 events)/(effective denominator sum). [Number of events increased to 2.3 for a 90% confidence level.]					
111000	0	<sup>10</sup> EBENHOH 74	HBC	$K^-p$ at rest	
105000	0	<sup>10</sup> SECHI-ZORN 73	HBC	$K^-p$ at rest	
<sup>10</sup> Effective denominator calculated by us.					

$\Gamma(n\mu^+\nu_\mu)/\Gamma(n\pi^+)$				$\Gamma_7/\Gamma_2$
Test of $\Delta S = \Delta Q$ rule.				
<u>EFFECTIVE DENOM.</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	
<b>&lt; 6.2 x 10<sup>-5</sup> OUR LIMIT</b> Our 90% CL limit = (6.7 events)/(effective denominator sum). [Number of events increased to 6.7 for a 90% confidence level.]				
33800	0	BAGGETT	69B	HBC
62000	2	<sup>11</sup> EISELE	69B	HBC
10150	0	<sup>12</sup> COURANT	64	HBC
1710	0	<sup>12</sup> NAUENBERG	64	HBC
120	1	GALTIERI	62	EMUL

<sup>11</sup> Effective denominator calculated by us.  
<sup>12</sup> Effective denominator taken from EISELE 67.

$\Gamma(pe^+ e^-)/\Gamma_{\text{total}}$		$\Gamma_8/\Gamma$	
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN COMMENT	
<b>&lt;7</b>			
	<sup>13</sup> ANG 69B	69B HBC	$K^- p$ at rest

<sup>13</sup> ANG 69B found three  $pe^+ e^-$  events in agreement with  $\gamma \rightarrow e^+ e^-$  conversion from  $\Sigma^+ \rightarrow p\gamma$ . The limit given here is for neutral currents.

$\Gamma(p\mu^+\mu^-)/\Gamma_{\text{total}}$			$\Gamma_9/\Gamma$	
A test for a $\Delta S = 1$ weak neutral current, but also allowed by higher-order electroweak interactions.				
VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>8.6^{+6.6}_{-5.4} \pm 5.5</math></b>	3	<sup>14</sup> PARK	05	HYCP $p$ Cu, 800 GeV

<sup>14</sup> The masses of the three dimuons of PARK 05 are within 1 MeV of one another, perhaps indicating the existence of a new state  $P^0$  with mass  $214.3 \pm 0.5$  MeV. In that case, the decay is  $\Sigma^+ \rightarrow pP^0$ ,  $P^0 \rightarrow \mu^+ \mu^-$ , with a branching fraction of  $(3.1^{+2.4}_{-1.9} \pm 1.5) \times 10^{-8}$ .

$\Gamma(\Sigma^+ \rightarrow ne^+ \nu_e)/\Gamma(\Sigma^- \rightarrow ne^- \bar{\nu}_e)$		$\Gamma_6/\Gamma_3^{\Sigma^-}$	
VALUE	CL% EVTS	DOCUMENT ID	TECN COMMENT
<b>&lt;0.009 OUR LIMIT</b> Our 90% CL limit, using $\Gamma(ne^+ \nu_e)/\Gamma(n\pi^+)$ above.			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.019	90	0	EBENHOH 74 HBC $K^- p$ at rest
<0.018	90	0	SECHI-ZORN 73 HBC $K^- p$ at rest
<0.12	95	0	COLE 71 HBC $K^- p$ at rest
<0.03	90	0	EISELE 69B HBC See EBENHOH 74

$\Gamma(\Sigma^+ \rightarrow n\mu^+ \nu_\mu)/\Gamma(\Sigma^- \rightarrow n\mu^- \bar{\nu}_\mu)$		$\Gamma_7/\Gamma_4^{\Sigma^-}$	
VALUE	CL% EVTS	DOCUMENT ID	TECN COMMENT
<b>&lt;0.12 OUR LIMIT</b> Our 90% CL limit, using $\Gamma(n\mu^+ \nu_\mu)/\Gamma(n\pi^+)$ above.			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.06+ <sup>0.045</sup> <sub>−0.03</sub>	2	EISELE 69B	HBC $K^- p$ at rest

$\Gamma(\Sigma^+ \rightarrow n\ell^+\nu)/\Gamma(\Sigma^- \rightarrow n\ell^-\bar{\nu})$		$(\Gamma_6+\Gamma_7)/(\Gamma_3^{\Sigma^-}+\Gamma_4^{\Sigma^-})$	
Test of $\Delta S = \Delta Q$ rule.			
VALUE	EVTS	DOCUMENT ID	TECN
<b>&lt;0.043 OUR LIMIT</b> Our 90% CL limit, using $[\Gamma(n e^+ \nu_e) + \Gamma(n \mu^+ \nu_\mu)]/\Gamma(n \pi^+)$ .			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.08	1	NORTON	69 HBC
<0.034	0	BAGGETT	67 HBC

 $\Sigma^+$  DECAY PARAMETERS

See the “Note on Baryon Decay Parameters” in the neutron Listings. A few early results have been omitted.

$\alpha_0$ FOR $\Sigma^+ \rightarrow p\pi^0$		$\alpha_0$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>−0.980+<sup>0.017</sup><sub>−0.015</sub> OUR FIT</b>			
<b>−0.980+<sup>0.017</sup><sub>−0.013</sub> OUR AVERAGE</b>			
−0.945+ <sup>0.055</sup> <sub>−0.042</sub>	1259	<sup>15</sup> LIPMAN 73	OSPK $\pi^+ p \rightarrow \Sigma^+$
−0.940±0.045	16k	BELLA MY 72	ASPK $\pi^+ p \rightarrow \Sigma^+ K^+$
−0.98 + <sup>0.05</sup> <sub>−0.02</sub>	1335	<sup>16</sup> HARRIS 70	OSPK $\pi^+ p \rightarrow \Sigma^+ K^+$
−0.999±0.022	32k	BANGERTER 69	HBC $K^- p$ 0.4 GeV/c

<sup>15</sup> Decay protons scattered off aluminum.<sup>16</sup> Decay protons scattered off carbon.

$\phi_0$ ANGLE FOR $\Sigma^+ \rightarrow p\pi^0$		$(\tan \phi_0 = \beta/\gamma)$	
VALUE (°)	EVTS	DOCUMENT ID	TECN COMMENT
<b>36 ±34 OUR AVERAGE</b>			
38.1+ <sup>35.7</sup> <sub>−37.1</sub>	1259	<sup>17</sup> LIPMAN 73	OSPK $\pi^+ p \rightarrow \Sigma^+ K^+$
22 ±90		<sup>18</sup> HARRIS 70	OSPK $\pi^+ p \rightarrow \Sigma^+ K^+$

<sup>17</sup> Decay proton scattered off aluminum.<sup>18</sup> Decay protons scattered off carbon.

$\alpha_+ / \alpha_0$				
Older results have been omitted.				
VALUE	<u>EVTs</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>−0.069±0.013 OUR FIT</b>				
<b>−0.073±0.021</b>	23k	MARRAFFINO 80	HBC	$K^- p$ 0.42–0.5 GeV/c

$\alpha_+$ FOR $\Sigma^+ \rightarrow n\pi^+$		$\alpha_+$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.068±0.013 OUR FIT</b>			
<b>0.066±0.016 OUR AVERAGE</b>			
0.037±0.049	4101	BERLEY 70B	HBC
0.069±0.017	35k	BANGERTER 69	HBC $K^- p$ 0.4 GeV/c

$\phi_+$ ANGLE FOR $\Sigma^+ \rightarrow n\pi^+$		$(\tan \phi_+ = \beta/\gamma)$	
VALUE (°)	EVTS	DOCUMENT ID	TECN COMMENT
<b>167±20 OUR AVERAGE</b> Error includes scale factor of 1.1.			
184±24	1054	<sup>19</sup> BERLEY 70B	HBC
143±29	560	BANGERTER 69B	HBC $K^- p$ 0.4 GeV/c

<sup>19</sup> Changed from 176 to 184° to agree with our sign convention.

$\alpha_\gamma$ FOR $\Sigma^+ \rightarrow p\gamma$		$\alpha_\gamma$	
VALUE	CL% EVTS	DOCUMENT ID	TECN COMMENT
<b>−0.76 ±0.08 OUR AVERAGE</b>			
−0.720±0.086±0.045	35k	<sup>20</sup> FOUCHER 92	SPEC $\Sigma^+$ 375 GeV
−0.86 ±0.13 ±0.04	190	KOBAYASHI 87	CNTR $\pi^+ p \rightarrow \Sigma^+ K^+$
−0.53 + <sup>0.38</sup> <sub>−0.36</sub>	46	MANZ 80	HBC $K^- p \rightarrow \Sigma^+ \pi^-$
−1.03 + <sup>0.52</sup> <sub>−0.42</sub>	61	GERSHWIN 69B	HBC $K^- p \rightarrow \Sigma^+ \pi^-$

<sup>20</sup> See TIMM 95 for a detailed description of the analysis.

# Baryon Particle Listings

$\Sigma^+$ ,  $\Sigma^0$ ,  $\Sigma^-$

## $\Sigma^+$ REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

PARK	05	PRL 94 021801	H.K. Park <i>et al.</i>	(FNAL HyperCP Collab.)
BARBOSA	00	PR D61 031101	R.F. Barbosa <i>et al.</i>	(FNAL E761 Collab.)
TIMM	95	PR D51 4638	S. Timm <i>et al.</i>	(FNAL E761 Collab.)
MORELOS	93	PRL 71 3417	A. Morelos <i>et al.</i>	(FNAL E761 Collab.)
FOUCHER	92	PRL 68 3004	M. Foucher <i>et al.</i>	(FNAL E761 Collab.)
HESSEY	89	ZPHY C42 175	N.P. Hessey <i>et al.</i>	(BNL-811 Collab.)
KOBAYASHI	87	PRL 59 868	N. Kobayashi <i>et al.</i>	(KYOT)
WILKINSON	87	PRL 58 855	C.A. Wilkinson <i>et al.</i>	(WISC, MICH, RUTG+)
BIAGI	85	ZPHY C28 495	S.F. Biagi <i>et al.</i>	(CERN WA62 Collab.)
ANKENBRA...	83	PRL 51 863	C.M. Ankenbrandt <i>et al.</i>	(FNAL, IOWA, ISU+)
MANZ	80	PL 96B 217	A. Manz <i>et al.</i>	(MPIM, VAND)
MARRAFFINO	80	PR D21 2501	J. Marraffino <i>et al.</i>	(VAND, MPIM)
NOWAK	78	NP B139 61	R.J. Nowak <i>et al.</i>	(LOUC, BELG, DURH+)
CONFORTO	76	NP B105 189	B. Conforto <i>et al.</i>	(RHEL, LOIC)
EBENHOH	74	ZPHY 266 367	H. Ebenhoh <i>et al.</i>	(HEIDT)
EBENHOH	73	ZPHY 264 413	W. Ebenhoh <i>et al.</i>	(HEIDT)
LIPMAN	73	PL 43B 89	N.H. Lipman <i>et al.</i>	(RHEL, SUSS, LOWC)
PDG	73	RMP 45 51	T.A. Lasinski <i>et al.</i>	(LBL, BRAN, CERN+)
SECHI-ZORN	73	PR D8 12	B. Sechi-Zorn, G.A. Snow	(UMD)
BELLAMY	72	PL 39B 299	E.H. Bellamy <i>et al.</i>	(LOWC, RHEL, SUSS)
BOHM	72	NP B48 1	G. Bohm <i>et al.</i>	(BERL, KIDR, BRUX, IASD, DUUC+)
Also		IJHE-73.2 Nov	G. Bohm	(BERL, KIDR, BRUX, IASD, DUUC+)
COLE	71	PR D4 631	J. Cole <i>et al.</i>	(STON, COLU)
TOVEE	71	NP B33 493	D.W. Tovee <i>et al.</i>	(LOUC, KIDR, BERL+)
BERLEY	70B	PR D1 2015	D. Berley <i>et al.</i>	(BNL, MASA, YALE)
EISELE	70	ZPHY 238 372	F. Eisele <i>et al.</i>	(HEID)
HARRIS	70	PRL 24 165	F. Harris <i>et al.</i>	(MICH, WISC)
PDG	70	RMP 42 87	A. Barbaro-Galtieri <i>et al.</i>	(LRL, BRAN+)
ANG	69B	ZPHY 228 151	G. Ang <i>et al.</i>	(HEID)
BAGGETT	69B	Thesis MDDP-TR-973	N.V. Baggett	(UMD)
BALTAY	69	PRL 22 615	C. Baltay <i>et al.</i>	(COLU, STON)
BANGERTER	69	Thesis UCLR 19244	R.O. Bangertter	(LRL)
BANGERTER	69B	PR 187 1821	R.O. Bangertter <i>et al.</i>	(LRL)
BARLOUTAUD	69	NP B14 153	R. Barloutaud <i>et al.</i>	(SACL, CERN, HEID)
EISELE	69	ZPHY 221 1	F. Eisele <i>et al.</i>	(HEID)
Also		PRL 13 291	W. Willis <i>et al.</i>	(BNL, CERN, HEID, UMD)
EISELE	69B	ZPHY 221 401	F. Eisele <i>et al.</i>	(HEID)
GERSHWIN	69B	PR 188 2077	L.K. Gershwin <i>et al.</i>	(LRL)
Also		Thesis UCLR 19246	L.K. Gershwin	(LRL)
NORTON	69	Thesis Nevis 175	H. Norton	(COLU)
BAGGETT	67	PRL 19 1458	N. Baggett <i>et al.</i>	(UMD)
Also		Vienna Abs. 374	N.V. Baggett, B. Kehoe	(UMD)
Also		Private Comm.	N.V. Baggett	(UMD)
BARASH	67	PRL 19 181	N. Barash <i>et al.</i>	(UMD)
EISELE	67	ZPHY 205 409	F. Eisele <i>et al.</i>	(HEID)
HYMAN	67	PL 25B 376	L.G. Hyman <i>et al.</i>	(ANL, CMU, NWES)
PDG	67	RMP 39 1	A.H. Rosenfeld <i>et al.</i>	(LRL, CERN, YALE)
CHANG	66	PR 151 1081	C.Y. Chang	(COLU)
Also		Thesis Nevis 145	C.Y. Chang	(COLU)
BAZIN	65	PRL 14 154	M. Bazin <i>et al.</i>	(PRIN, COLU)
BAZIN	65B	PR 140B 1358	M. Bazin <i>et al.</i>	(PRIN, RUTG, COLU)
SCHMIDT	65	PR 140B 1328	P. Schmidt	(COLU)
BHOWMIK	64	NP 53 22	B. Bhowmik <i>et al.</i>	(DELH)
COURANT	64	PR 136 B1791	H. Courant <i>et al.</i>	(CERN, HEID, UMD+)
NAUENBERG	64	PRL 12 679	U. Nauenberg <i>et al.</i>	(COLU, RUTG, PRIN)
BARKAS	63	PRL 11 26	W.H. Barkas, J.N. Dyer, H.H. Heckman	(LRL)
Also		Thesis UCLR 9450	J.N. Dyer	(LRL)
GALTIERI	62	PRL 9 26	A. Barbaro-Galtieri <i>et al.</i>	(LRL)
HUMPHREY	62	PR 127 1305	W.E. Humphrey, R.R. Ross	(LRL)

$\Sigma^0$

$I(J^P) = 1(\frac{1}{2}^+)$  Status: \* \* \* \*

COURANT 63 and ALFF 65, using  $\Sigma^0 \rightarrow \Lambda e^+ e^-$  decays (Dalitz decays), determined the  $\Sigma^0$  parity to be positive, given that  $J = 1/2$  and that certain very reasonable assumptions about form factors are true. The results of experiments involving the Primakoff effect, from which the  $\Sigma^0$  mean life and  $\Sigma^0 \rightarrow \Lambda$  transition magnetic moment come (see below), strongly support  $J = 1/2$ .

## $\Sigma^0$ MASS

The fit uses  $\Sigma^+$ ,  $\Sigma^0$ ,  $\Sigma^-$ , and  $\Lambda$  mass and mass-difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1192.642±0.024 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

1192.65	±0.020±0.014	3327	<sup>1</sup> WANG	97	SPEC	$\Sigma^0 \rightarrow \Lambda \gamma \rightarrow (p\pi^-)(e^+ e^-)$
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<sup>1</sup> This WANG 97 result is redundant with the  $\Sigma^0$ - $\Lambda$  mass-difference measurement below.

## $m_{\Sigma^-} - m_{\Sigma^0}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.807±0.035 OUR FIT</b>				Error includes scale factor of 1.1.
<b>4.86 ±0.08 OUR AVERAGE</b>				Error includes scale factor of 1.2.
4.87 ±0.12	37	DOSCH	65	HBC
5.01 ±0.12	12	SCHMIDT	65	HBC See note with $\Lambda$ mass
4.75 ±0.1	18	BURNSTEIN	64	HBC

## $m_{\Sigma^0} - m_{\Lambda}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>76.959±0.023 OUR FIT</b>				
<b>76.966±0.020±0.013</b>	3327	WANG	97	SPEC $\Sigma^0 \rightarrow \Lambda \gamma \rightarrow (p\pi^-)(e^+ e^-)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

76.23 ±0.55	109	COLAS	75	HLBC	$\Sigma^0 \rightarrow \Lambda \gamma$
76.63 ±0.28	208	SCHMIDT	65	HBC	See note with $\Lambda$ mass

## $\Sigma^0$ MEAN LIFE

These lifetimes are deduced from measurements of the cross sections for the Primakoff process  $\Lambda \rightarrow \Sigma^0$  in nuclear Coulomb fields. An alternative expression of the same information is the  $\Sigma^0$ - $\Lambda$  transition magnetic moment given in the following section. The relation is  $(\mu_{\Sigma\Lambda}/\mu_N)^2 \tau = 1.92951 \times 10^{-19}$  s (see DEVLIN 86).

VALUE ( $10^{-20}$ s)	DOCUMENT ID	TECN	COMMENT
<b>7.4±0.7 OUR EVALUATION</b>	Using $\mu_{\Sigma\Lambda}$ (see the above note).		
6.5+1.7 -1.1	<sup>2</sup> DEVLIN	86	SPEC Primakoff effect
7.6±0.5±0.7	<sup>3</sup> PETERSEN	86	SPEC Primakoff effect
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5.8±1.3	<sup>2</sup> DYDAK	77	SPEC See DEVLIN 86
<sup>2</sup> DEVLIN 86 is a recalculation of the results of DYDAK 77 removing a numerical approximation made in that work.			
<sup>3</sup> An additional uncertainty of the Primakoff formalism is estimated to be < 5%.			

## $|\mu(\Sigma^0 \rightarrow \Lambda)|$ TRANSITION MAGNETIC MOMENT

See the note in the  $\Sigma^0$  mean-life section above. Also, see the “Note on Baryon Magnetic Moments” in the  $\Lambda$  Listings.

VALUE ( $\mu_N$ )	DOCUMENT ID	TECN	COMMENT
<b>1.61±0.08 OUR AVERAGE</b>			
1.72+0.17 -0.19	<sup>4</sup> DEVLIN	86	SPEC Primakoff effect
1.59±0.05±0.07	<sup>5</sup> PETERSEN	86	SPEC Primakoff effect
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.82+0.25 -0.18	<sup>4</sup> DYDAK	77	SPEC See DEVLIN 86
<sup>4</sup> DEVLIN 86 is a recalculation of the results of DYDAK 77 removing a numerical approximation made in that work.			
<sup>5</sup> An additional uncertainty of the Primakoff formalism is estimated to be < 2.5%.			

## $\Sigma^0$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \quad \Lambda \gamma$	100 %	
$\Gamma_2 \quad \Lambda \gamma \gamma$	< 3 %	90%
$\Gamma_3 \quad \Lambda e^+ e^-$	[a] $5 \times 10^{-3}$	

[a] A theoretical value using QED.

## $\Sigma^0$ BRANCHING RATIOS

$\Gamma(\Lambda \gamma \gamma)/\Gamma_{\text{total}}$	CL%	DOCUMENT ID	TECN	$\Gamma_2/\Gamma$
<b>&lt;0.03</b>	90	COLAS	75	HLBC

$\Gamma(\Lambda e^+ e^-)/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
See COURANT 63 and ALFF 65 for measurements of the invariant-mass spectrum of the Dalitz pairs.	

VALUE	DOCUMENT ID	COMMENT
<b>0.00545</b>	FEINBERG	58 Theoretical QED calculation

## $\Sigma^0$ REFERENCES

WANG	97	PR D56 2544	M.H.L.S. Wang <i>et al.</i>	(BNL-E766 Collab.)
DEVLIN	86	PR D34 1626	T. Devlin, P.C. Petersen, A. Beretvas	(RUTG)
PETERSEN	86	PRL 57 949	P.C. Petersen <i>et al.</i>	(RUTG, WISC, MICH+)
DYDAK	77	NP B118 1	F. Dydak <i>et al.</i>	(CERN, DORT, HEIDH)
COLAS	75	NP B91 253	J. Colas <i>et al.</i>	(ORSAY)
ALFF	65	PR 137 B1105	C. Alff <i>et al.</i>	(COLU, RUTG, BNL) <sup>P</sup>
DOSCH	65	PL 14 239	H.C. Dosch <i>et al.</i>	(HEID)
SCHMIDT	65	PR 140B 1328	P. Schmidt	(COLU)
BURNSTEIN	64	PRL 13 66	R.A. Burnstein <i>et al.</i>	(UMD)
COURANT	63	PRL 10 409	H. Courant <i>et al.</i>	(CERN, UMD) <sup>P</sup>
FEINBERG	58	PR 109 1019	G. Feinberg	(BNL)

$\Sigma^-$

$I(J^P) = 1(\frac{1}{2}^+)$  Status: \* \* \* \*

We have omitted some results that have been superseded by later experiments. See our earlier editions.

## $\Sigma^-$ MASS

The fit uses  $\Sigma^+$ ,  $\Sigma^0$ ,  $\Sigma^-$ , and  $\Lambda$  mass and mass-difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1197.449 ± 0.030 OUR FIT</b>				Error includes scale factor of 1.2.
<b>1197.45 ± 0.04 OUR AVERAGE</b>				Error includes scale factor of 1.2.
1197.417 ± 0.040		GUREV 93	SPEC	$\Sigma^-$ -C atom, crystal diff.
1197.532 ± 0.057		GALL 88	CNTR	$\Sigma^-$ -Pb, $\Sigma^-$ -W atoms
1197.43 ± 0.08	3000	SCHMIDT 65	HBC	See note with $\Lambda$ mass
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1197.24 ± 0.15		<sup>1</sup> DUGAN 75	CNTR	Exotic atoms
<sup>1</sup> GALL 88 concludes that the DUGAN 75 mass needs to be reevaluated.				

$m_{\Sigma^-} - m_{\Sigma^+}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.08 ± 0.08 OUR FIT</b>				Error includes scale factor of 1.9.
<b>8.09 ± 0.16 OUR AVERAGE</b>				
7.91 ± 0.23	86	BOHM 72	EMUL	
8.25 ± 0.25	2500	DOSCH 65	HBC	
8.25 ± 0.40	87	BARKAS 63	EMUL	

$m_{\Sigma^-} - m_{\Lambda}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>81.766 ± 0.030 OUR FIT</b>				Error includes scale factor of 1.2.
<b>81.69 ± 0.07 OUR AVERAGE</b>				
81.64 ± 0.09	2279	HEPP 68	HBC	
81.80 ± 0.13	85	SCHMIDT 65	HBC	See note with $\Lambda$ mass
81.70 ± 0.19		BURNSTEIN 64	HBC	

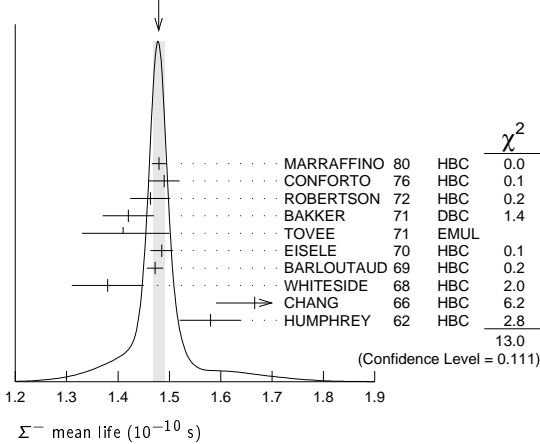
$\Sigma^-$  MEAN LIFE

Measurements with an error  $\geq 0.2 \times 10^{-10}$  s have been omitted.

VALUE ( $10^{-10}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.479 ± 0.011 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
1.480 ± 0.014	16k	MARRAFFINO 80	HBC	$K^- p$ 0.42–0.5 GeV/c
1.49 ± 0.03	8437	CONFORTO 76	HBC	$K^- p$ 1–1.4 GeV/c
1.463 ± 0.039	2400	ROBERTSON 72	HBC	$K^- p$ 0.25 GeV/c
1.42 ± 0.05	1383	BAKKER 71	DBC	$K^- N \rightarrow \Sigma^- \pi \pi$
1.41 ± 0.09		TOVEE 71	EMUL	
1.485 ± 0.022	100k	EISELE 70	HBC	$K^- p$ at rest
1.472 ± 0.016	10k	BARLOUTAUD 69	HBC	$K^- p$ 0.4–1.2 GeV/c
1.38 ± 0.07	506	WHITESIDE 68	HBC	$K^- p$ at rest
1.666 ± 0.075	3267	<sup>2</sup> CHANG 66	HBC	$K^- p$ at rest
1.58 ± 0.06	1208	HUMPHREY 62	HBC	$K^- p$ at rest

<sup>2</sup> We have increased the CHANG 66 error of 0.026; see our 1970 edition, Reviews of Modern Physics **42** 87 (1970).

WEIGHTED AVERAGE  
1.479 ± 0.011 (Error scaled by 1.3)

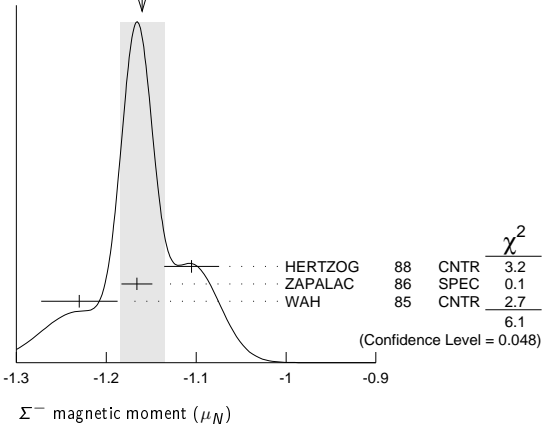


$\Sigma^-$  MAGNETIC MOMENT

See the "Note on Baryon Magnetic Moments" in the  $\Lambda$  Listings. Measurements with an error  $\geq 0.3 \mu_N$  have been omitted.

VALUE ( $\mu_N$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-1.160 ± 0.025 OUR AVERAGE</b>				Error includes scale factor of 1.7. See the ideogram below.
-1.105 ± 0.029 ± 0.010		HERTZOG 88	CNTR	$\Sigma^-$ -Pb, $\Sigma^-$ -W atoms
-1.166 ± 0.014 ± 0.010	671k	ZAPALAC 86	SPEC	$ne^- \nu, \pi \pi^-$ decays
-1.23 ± 0.03 ± 0.03		WAH 85	CNTR	$p\text{Cu} \rightarrow \Sigma^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.89 ± 0.14	516k	DECK 83	SPEC	$p\text{Be} \rightarrow \Sigma^- X$

WEIGHTED AVERAGE  
-1.160 ± 0.025 (Error scaled by 1.7)



$\Sigma^-$  CHARGE RADIUS

VALUE (fm)	DOCUMENT ID	TECN	COMMENT
<b>0.780 ± 0.080 ± 0.060</b>	<sup>3</sup> ESCHRICH 01	SELX	$\Sigma^- e \rightarrow \Sigma^- e$
<sup>3</sup> ESCHRICH 01 actually gives $\langle r^2 \rangle = (0.61 \pm 0.12 \pm 0.09) \text{ fm}^2$ .			

$\Sigma^-$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad n \pi^-$	(99.848 ± 0.005) %
$\Gamma_2 \quad n \pi^- \gamma$	[a] ( 4.6 ± 0.6 ) $\times 10^{-4}$
$\Gamma_3 \quad n e^- \bar{\nu}_e$	( 1.017 ± 0.034 ) $\times 10^{-3}$
$\Gamma_4 \quad n \mu^- \bar{\nu}_\mu$	( 4.5 ± 0.4 ) $\times 10^{-4}$
$\Gamma_5 \quad \Lambda e^- \bar{\nu}_e$	( 5.73 ± 0.27 ) $\times 10^{-5}$

[a] See the Listings below for the pion momentum range used in this measurement.

CONSTRAINED FIT INFORMATION

An overall fit to 3 branching ratios uses 16 measurements and one constraint to determine 4 parameters. The overall fit has a  $\chi^2 = 8.7$  for 13 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	-64		
$x_4$	-77	0	
$x_5$	-5	0	0
	$x_1$	$x_3$	$x_4$

$\Sigma^-$  BRANCHING RATIOS

$\Gamma(n \pi^- \gamma)/\Gamma(n \pi^-)$   $\Gamma_2/\Gamma_1$   
The  $\pi^+$  momentum cuts differ, so we do not average the results but simply use the latest value for the Summary Table.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.46 ± 0.06</b>	292	EBENHOH 73	HBC	$\pi^+ < 150 \text{ MeV/c}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.10 ± 0.02	23	ANG 69B	HBC	$\pi^+ < 110 \text{ MeV/c}$
~ 1.1		BAZIN 65B	HBC	$\pi^+ < 166 \text{ MeV/c}$

$\Gamma(ne^- \bar{\nu}_e)/\Gamma(n \pi^-)$   $\Gamma_3/\Gamma_1$

Measurements with an error  $\geq 0.2 \times 10^{-3}$  have been omitted.

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.019 ± 0.035 OUR FIT</b>				
<b>1.019 ± 0.031 - 0.040 OUR AVERAGE</b>				
0.96 ± 0.05	2847	BOURQUIN 83c	SPEC	SPS hyperon beam
1.09 ± 0.06 - 0.08	601	<sup>4</sup> EBENHOH 74	HBC	$K^- p$ at rest
1.05 ± 0.07 - 0.13	455	<sup>4</sup> SECHI-ZORN 73	HBC	$K^- p$ at rest
0.97 ± 0.15	57	COLE 71	HBC	$K^- p$ at rest
1.11 ± 0.09	180	BIERMAN 68	HBC	

<sup>4</sup> An additional negative systematic error is included for internal radiative corrections and latest form factors; see BOURQUIN 83c.

# Baryon Particle Listings

## $\Sigma^-$

$\Gamma(n\mu^- \bar{\nu}_\mu)/\Gamma(n\pi^-)$		$\Gamma_4/\Gamma_1$			
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.45±0.04 OUR FIT</b>					
<b>0.45±0.04 OUR AVERAGE</b>					
0.38±0.11	13	COLE	71	HBC	$K^-p$ at rest
0.43±0.06	72	ANG	69	HBC	$K^-p$ at rest
0.43±0.09	56	BAGGETT	69	HBC	$K^-p$ at rest
0.56±0.20	11	BAZIN	65B	HBC	$K^-p$ at rest
0.66±0.15	22	COURANT	64	HBC	

$\Gamma(\Lambda e^- \bar{\nu}_e)/\Gamma(n\pi^-)$			$\Gamma_5/\Gamma_1$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>0.574±0.027 OUR FIT</b>						
<b>0.574±0.027 OUR AVERAGE</b>						
0.561±0.031	1620	5	BOURQUIN	82	SPEC	SPS hyperon beam
0.63 ±0.11	114		THOMPSON	80	ASPK	Hyperon beam
0.52±0.09	31		BALTAY	69	HBC	$K^-p$ at rest
0.69 ±0.12	31		EISELE	69	HBC	$K^-p$ at rest
0.64±0.12	35		BARASH	67	HBC	$K^-p$ at rest
0.75 ±0.28	11		COURANT	64	HBC	$K^-p$ at rest

<sup>5</sup> The value is from BOURQUIN 83B, and includes radiation corrections and new acceptance.

### $\Sigma^-$ DECAY PARAMETERS

See the “Note on Baryon Decay Parameters” in the neutron Listings.  
Older, outdated results have been omitted.

#### $\alpha_-$ FOR $\Sigma^- \rightarrow n\pi^-$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>−0.068 ± 0.008 OUR AVERAGE</b>				
−0.062 ± 0.024	28k	HANSL	78	HBC $K^-p \rightarrow \Sigma^- \pi^+$
−0.067 ± 0.011	60k	BOGERT	70	HBC $K^-p$ 0.4 GeV/c
−0.071 ± 0.012	51k	BANGERTER	69	HBC $K^-p$ 0.4 GeV/c

#### $\phi$ ANGLE FOR $\Sigma^- \rightarrow n\pi^-$

( $\tan\phi = \beta/\gamma$ )

VALUE (°)	EVTS	DOCUMENT ID	TECN	COMMENT
10±15 OUR AVERAGE				
+ 5±23	1092	6 BERLEY	70B HBC	<i>n</i> rescattering
14±19	1385	BANGERTER	69B HBC	<i>K</i> <sup>-</sup> <i>p</i> 0.4 GeV/ <i>c</i>

<sup>6</sup> BERLEY 70B changed from −5 to +5° to agree with our sign convention.

#### $g_A/g_V$ FOR $\Sigma^- \rightarrow ne^- \bar{\nu}_e$

Measurements with fewer than 500 events have been omitted. Where necessary, signs have been changed to agree with our conventions, which are given in the “Note on Baryon Decay Parameters” in the neutron Listings. What is actually listed is  $|g_1/f_1 - 0.237g_2/f_1|$ . This reduces to  $g_A/g_V \equiv g_1(0)/f_1(0)$  on making the usual assumption that  $g_2 = 0$ . See also the note on HSUEH 88.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.340±0.017 OUR AVERAGE				
+0.327±0.007±0.019	50k	7 HSUEH	88	SPEC $\Sigma^-$ 250 GeV
+0.34 ±0.05	4456	8 BOURQUIN	83c	SPEC SPS hyperon beam
0.385±0.037	3507	9 TANENBAUM	74	ASPK
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.29 ±0.07	25k	HSUEH	85	SPEC See HSUEH 88
0.17 <sup>+0.07</sup> <sub>-0.09</sub>	519	DECAMP	77	ELEC Hyperon beam

<sup>7</sup> The sign is, with our conventions, unambiguously positive. The value assumes, as usual, that  $g_2 = 0$ . If  $g_2$  is included in the fit, than (with our sign convention)  $g_2 = -0.56 \pm 0.37$ , with a corresponding reduction of  $g_A/g_V$  to  $+0.20 \pm 0.08$ .

<sup>8</sup> BOURQUIN 83c favors the positive sign by at least 2.6 standard deviations.

<sup>9</sup> TANENBAUM 74 gives  $0.435 \pm 0.035$ , assuming no  $q^2$  dependence in  $g_A$  and  $g_V$ . The listed result allows  $q^2$  dependence, and is taken from HSUEH 88.

#### $f_2(0)/f_1(0)$ FOR $\Sigma^- \rightarrow ne^- \bar{\nu}_e$

The signs have been changed to be in accord with our conventions, given in the “Note on Baryon Decay Parameters” in the neutron Listings.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.97±0.14 OUR AVERAGE				
+ 0.96±0.07±0.13	50k	HSUEH	88	SPEC Σ <sup>-</sup> 250 GeV
+ 1.02±0.34	4456	BOURQUIN	83c	SPEC SPS hyperon beam

#### TRIPLE CORRELATION COEFFICIENT $D$ for $\Sigma^- \rightarrow ne^- \bar{\nu}_e$

The coefficient  $D$  of the term  $D \mathbf{P} \cdot (\mathbf{p}_e \times \mathbf{p}_{\nu})$  in the  $\Sigma^- \rightarrow ne^- \bar{\nu}_e$  decay angular distribution. A nonzero value would indicate a violation of time-reversal invariance.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.11±0.10	50k	HSUEH	88	SPEC
				Σ <sup>-</sup> 250 GeV

#### $g_V/g_A$ FOR $\Sigma^- \rightarrow \Lambda e^- \bar{\nu}_e$

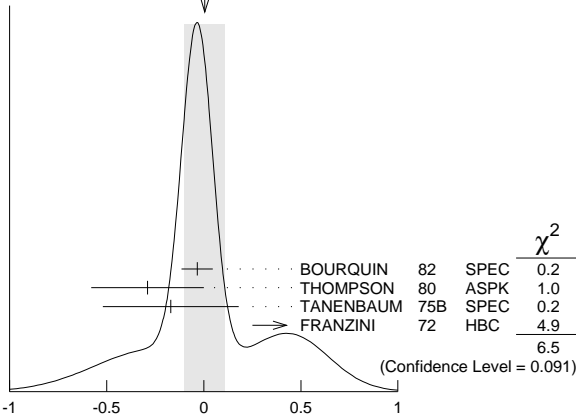
For the sign convention, see the “Note on Baryon Decay Parameters” in the neutron Listings. The value is predicted to be zero by conserved vector current theory. The values averaged assume CVC-SU(3) weak magnetism term.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.01 ±0.10 OUR AVERAGE</b>				
Error includes scale factor of 1.5. See the ideogram below.				
−0.034 ±0.080	1620	<sup>10</sup> BOURQUIN	82	SPEC SPS hyperon beam
−0.29 ±0.29	114	THOMPSON	80	ASPK BNL hyperon beam
−0.17 ±0.35	55	TANENBAUM	75B	SPEC BNL hyperon beam
+0.45 ±0.20	186	<sup>10,11</sup> FRANZINI	72	HBC

<sup>10</sup> The sign has been changed to agree with our convention.

<sup>11</sup> The FRANZINI 72 value includes the events of earlier papers.

WEIGHTED AVERAGE  
0.01±0.10 (Error scaled by 1.5)



#### $g_{WM}/g_A$ FOR $\Sigma^- \rightarrow \Lambda e^- \bar{\nu}_e$

The values quoted assume the CVC prediction  $g_V = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
2.4 ±1.7 OUR AVERAGE					
1.75 ±3.5	114	THOMPSON	80	ASPK	BNL hyperon beam
3.5 ±4.5	55	TANENBAUM	75B	SPEC	BNL hyperon beam
2.4 ±2.1	186	FRANZINI	72	HBC	

### $\Sigma^-$ REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

ESCHRICH	01	PL B522 233	I. Eschrich <i>et al.</i>	(FNAL SELEX Collab.)
GUREV	93	JETPL 57 400	M.P. Gurev <i>et al.</i>	(PNPI)
		Translated from ZETFP 57 389.		
GALL	88	PRL 60 186	K.P. Gall <i>et al.</i>	(BOST, MIT, WILL, CIT+)
HERTZOG	88	PR D37 1142	D.W. Hertzog <i>et al.</i>	(WILL, BOST, MIT+)
HSUEH	88	PR D38 2056	S.Y. Hsueh <i>et al.</i>	(CHIC, ELMT, FNAL+)
ZAPALAC	86	PRL 57 1526	G. Zapalac <i>et al.</i>	(EFI, ELMT, FNAL+)
HSUEH	85	PRL 54 2399	S.Y. Hsueh <i>et al.</i>	(CHIC, ELMT, FNAL+)
WAH	85	PRL 55 2551	Y.W. Wah <i>et al.</i>	(FNAL, IOWA, ISU)
BOURQUIN	83B	ZPHY C21 27	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
BOURQUIN	83C	ZPHY C21 17	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
DECK	83	PR D28 1	L. Deck <i>et al.</i>	(RUTG, WISC, MICH, MINN)
BOURQUIN	82	ZPHY C12 307	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
MARRAFFINO	80	PR D21 2501	J. Marraffino <i>et al.</i>	(VAND, MPIM)
THOMPSON	80	PR D21 25	J.A. Thompson <i>et al.</i>	(PITT, BNL)
HANSL	78	NP B132 45	T. Hansl <i>et al.</i>	(MPIM, VAND)
DECAMP	77	PL 66B 295	D. Decamp <i>et al.</i>	(LALO, EPOL)
CONFORTO	76	NP B105 189	B. Conforto <i>et al.</i>	(RHEL, LOIC)
DUGAN	75	NP A254 396	G. Dugan <i>et al.</i>	(COLU, YALE)
TANENBAUM	75B	PR D12 1871	W. Tanenbaum <i>et al.</i>	(YALE, FNAL, BNL)
EBENHOH	74	ZPHY 266 367	H. Ebenhoeh <i>et al.</i>	(HEIDT)
TANENBAUM	74	PRL 33 175	W. Tanenbaum <i>et al.</i>	(YALE, FNAL, BNL)
EBENHOH	73	ZPHY 264 413	W. Ebenhoeh <i>et al.</i>	(HEIDT)
SECHI-ZORN	73	PR D8 12	B. Sechi-Zorn, G.A. Snow	(UMD)
BOHM	72	NP B48 1	G. Bohm <i>et al.</i>	(BERL, KIDR, BRUX, IASD+)
FRANZINI	72	PR D6 2417	P. Franzini <i>et al.</i>	(COLU, HEID, UMD+)
ROBERTSON	72	Thesis UMI 78-00877	R.M. Robertson	(IIT)
BAKKER	71	LNC 1 37	A.M. Bakker <i>et al.</i>	(SABRE Collab.)
COLE	71	PR D4 631	J. Cole <i>et al.</i>	(STON, COLU)
Also		Thesis Nevis 175	H. Norton	(COLU)
TOVEE	71	NP B33 493	D.M. Tovee <i>et al.</i>	(LOUC, KIDR, BERL+)
BERLEY	70B	PR D1 2015	D. Berley <i>et al.</i>	(BNL, MASA, YALE)
BOGERT	70	PR D2 6	D.V. Bogert <i>et al.</i>	(BNL, MASA, YALE)
EISELE	70	ZPHY 238 372	F. Eisele <i>et al.</i>	(HEID)
PDG	70	RMP 42 87	A. Barbaro-Gatti <i>et al.</i>	(LRL, BRAN+)
ANG	69	ZPHY 223 103	G. Ang <i>et al.</i>	(HEID)
ANG	69B	ZPHY 228 151	G. Ang <i>et al.</i>	(HEID)
BAGGETT	69	PRL 23 249	N.V. Baggett, B. Kehoe, G.A. Snow	(UMD)
BALTAY	69	PRL 22 615	C. Baltay <i>et al.</i>	(COLU, STON)
BANGERTER	69	Thesis UCLRL 19244	R.O. Bangerter	(LRL)
BANGERTER	69B	PR 187 1821	R.O. Bangerter <i>et al.</i>	(LRL)
BARLOUTAUD	69	NP B14 153	R. Barloutaud <i>et al.</i>	(SACL, CERN, HEID)
EISELE	69	ZPHY 221 1	F. Eisele <i>et al.</i>	(HEID)
BIERMAN	68	PRL 20 1459	E. Bierman <i>et al.</i>	(PRIN)
HEPP	68	ZPHY 214 71	V. Hepp, H. Schleich	(HEID)
WHITESIDE	68	NC 54A 537	H. Whiteside, J. Gollub	(OBER)
BARASH	67	PRL 19 181	N. Barash <i>et al.</i>	(UMD)
CHANG	66	PR 151 1081	C.Y. Chang	(COLU)
BAZIN	65B	PR 140B 1358	M. Bazin <i>et al.</i>	(PRIN, RUTG, COLU)
DOSCH	65	PL 14 239	H.C. Dosch <i>et al.</i>	(HEID)
Also		PR 151 1081	C.Y. Chang	(COLU)
SCHMIDT	65	PR 140B 1328	P. Schmidt	(COLU)
BURNSTEIN	64	PRL 13 666	R.A. Burnstein <i>et al.</i>	(UMD)
COURANT	64	PR 136 B1791	L. Courant <i>et al.</i>	(CERN, HEID, UMD+)
BARKAS	63	PRL 11 26	W.H. Barkas, J.N. Dyer, H.H. Heckman	(LRL)
HUMPHREY	62	PR 127 1305	W.E. Humphrey, R.R. Ross	(LRL)

$\Sigma(1385)$  $\Sigma(1385) \ 3/2^+$ 

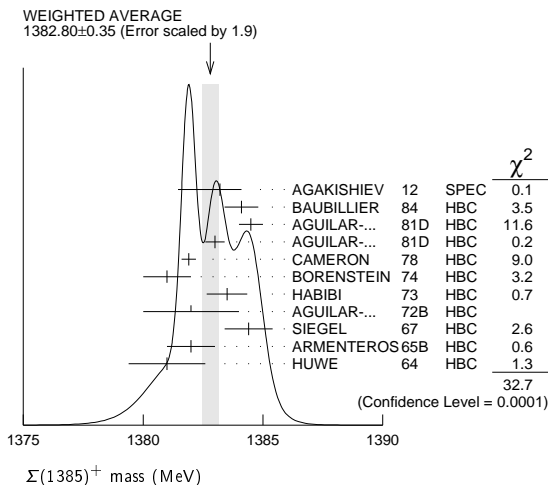
$$I(J^P) = 1(\frac{3}{2}^+) \text{ Status: } ****$$

Discovered by ALSTON 60. Early measurements of the mass and width for combined charge states have been omitted. They may be found in our 1984 edition Reviews of Modern Physics **56** S1 (1984).

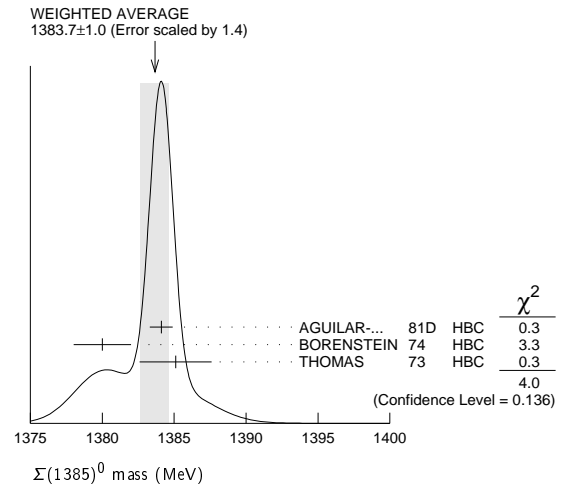
We average only the most significant determinations. We do not average results from inclusive experiments with large backgrounds or results which are not accompanied by some discussion of experimental resolution. Nevertheless systematic differences between experiments remain. (See the ideograms in the Listings below.) These differences could arise from interference effects that change with production mechanism and/or beam momentum. They can also be accounted for in part by differences in the parametrizations employed. (See BORENSTEIN 74 for a discussion on this point.) Thus BORENSTEIN 74 uses a Breit-Wigner with energy-independent width, since a  $P$ -wave was found to give unsatisfactory fits. CAMERON 78 uses the same form. On the other hand HOLMGREN 77 obtains a good fit to their  $\Lambda\pi$  spectrum with a  $P$ -wave Breit-Wigner, but includes the partial width for the  $\Sigma\pi$  decay mode in the parametrization. AGUILAR-BENITEZ 81D gives masses and widths for five different Breit-Wigner shapes. The results vary considerably. Only the best-fit  $S$ -wave results are given here.

 $\Sigma(1385) \text{ MASSES}$  $\Sigma(1385)^+ \text{ MASS}$ 

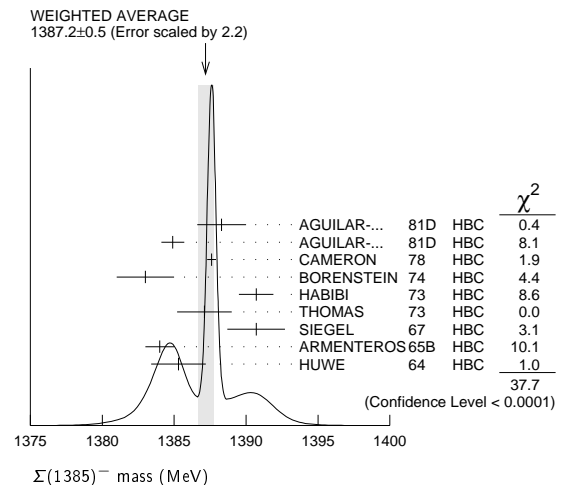
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1382.80 <math>\pm</math> 0.35 OUR AVERAGE</b>		Error includes scale factor of 1.9. See the ideogram below.		
1383.2 $\pm$ 0.9 $^{+0.1}_{-1.5}$		AGAKISHIEV 12	SPEC	$pp \rightarrow \Sigma(1385)^+ K^+ n$ , 3.5 GeV
1384.1 $\pm$ 0.7	1897	BAUBILLIER 84	HBC	$K^- p \rightarrow \Sigma(1385)^+ K^+ n$ , 8.25 GeV/c
1384.5 $\pm$ 0.5	5256	AGUILAR-... 81D	HBC	$K^- p \rightarrow \Lambda\pi\pi$ 4.2 GeV/c
1383.0 $\pm$ 0.4	9361	AGUILAR-... 81D	HBC	$K^- p \rightarrow \Lambda\pi\pi$ 4.2 GeV/c
1381.9 $\pm$ 0.3	6900	CAMERON 78	HBC	$K^- p$ 0.96-1.36 GeV/c
1381 $\pm$ 1	6846	BORENSTEIN 74	HBC	$K^- p$ 2.18 GeV/c
1383.5 $\pm$ 0.85	2300	HABIBI 73	HBC	$K^- p \rightarrow \Lambda\pi\pi$
1382 $\pm$ 2	400	AGUILAR-... 72B	HBC	$K^- p \rightarrow \Lambda\pi$ 's
1384.4 $\pm$ 1.0	1260	SIEGEL 67	HBC	$K^- p$ 2.1 GeV/c
1382 $\pm$ 1	750	ARMENTEROS65B	HBC	$K^- p$ 0.9-1.2 GeV/c
1381.0 $\pm$ 1.6	859	HUWE 64	HBC	$K^- p$ 1.22 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1385.1 $\pm$ 1.2	600	BAKER 80	HYBR	$\pi^+ p$ 7 GeV/c
1383.2 $\pm$ 1.0	750	BAKER 80	HYBR	$K^- p$ 7 GeV/c
1381 $\pm$ 2	7k	1 BAUBILLIER 79B	HBC	$K^- p$ 8.25 GeV/c
1391 $\pm$ 2	2k	CAUTIS 79	HYBR	$\pi^+ p/K^- p$ 11.5 GeV
1390 $\pm$ 2	100	1 SUGAHARA 79B	HBC	$\pi^- p$ 6 GeV/c
1385 $\pm$ 3	22k	1,2 BARREIRO 77B	HBC	$K^- p$ 4.2 GeV/c
1385 $\pm$ 1	2594	HOLMGREN 77	HBC	See AGUILAR-BENITEZ 81D
1380 $\pm$ 2		1 BARDADIN-... 75	HBC	$K^- p$ 14.3 GeV/c
1382 $\pm$ 1	3740	3 BERTHON 74	HBC	$K^- p$ 1263-1843 MeV/c
1390 $\pm$ 6	46	AGUILAR-... 70B	HBC	$K^- p \rightarrow \Sigma\pi$ 's 4 GeV/c
1383 $\pm$ 8	62	4 BIRMINGHAM 66	HBC	$K^- p$ 3.5 GeV/c
1378 $\pm$ 5	135	LONDON 66	HBC	$K^- p$ 2.24 GeV/c
1384.3 $\pm$ 1.9	250	4 SMITH 65	HBC	$K^- p$ 1.8 GeV/c
1382.6 $\pm$ 2.1	250	4 SMITH 65	HBC	$K^- p$ 1.95 GeV/c
1375.0 $\pm$ 3.9	170	COOPER 64	HBC	$K^- p$ 1.45 GeV/c
1376.0 $\pm$ 3.9	154	4 ELY 61	HLBC	$K^- p$ 1.11 GeV/c

 $\Sigma(1385)^0 \text{ MASS}$ 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1383.7 <math>\pm</math> 1.0 OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.		
1384.1 $\pm$ 0.8	5722	AGUILAR-... 81D	HBC	$K^- p \rightarrow \Lambda\pi\pi$ 4.2 GeV/c
1380 $\pm$ 2	3100	5 BORENSTEIN 74	HBC	$K^- p \rightarrow \Lambda\pi\pi$ 2.18 GeV/c
1385.1 $\pm$ 2.5	240	4 THOMAS 73	HBC	$\pi^- p \rightarrow \Lambda\pi^0 K^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1389 $\pm$ 3	500	6 BAUBILLIER 79B	HBC	$K^- p$ 8.25 GeV/c

 $\Sigma(1385)^- \text{ MASS}$ 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1387.2 <math>\pm</math> 0.5 OUR AVERAGE</b>		Error includes scale factor of 2.2. See the ideogram below.		
1388.3 $\pm$ 1.7	620	AGUILAR-... 81D	HBC	$K^- p \rightarrow \Lambda\pi\pi$ 4.2 GeV/c
1384.9 $\pm$ 0.8	3346	AGUILAR-... 81D	HBC	$K^- p \rightarrow \Lambda\pi\pi$ 4.2 GeV/c
1387.6 $\pm$ 0.3	9720	CAMERON 78	HBC	$K^- p$ 0.96-1.36 GeV/c
1383 $\pm$ 2	2303	BORENSTEIN 74	HBC	$K^- p$ 2.18 GeV/c
1390.7 $\pm$ 1.2	1900	HABIBI 73	HBC	$K^- p \rightarrow \Lambda\pi\pi$
1387.1 $\pm$ 1.9	630	4 THOMAS 73	HBC	$\pi^- p \rightarrow \Lambda\pi^- K^+$
1390.7 $\pm$ 2.0	370	SIEGEL 67	HBC	$K^- p$ 2.1 GeV/c
1384 $\pm$ 1	1380	ARMENTEROS65B	HBC	$K^- p$ 0.9-1.2 GeV/c
1385.3 $\pm$ 1.9	1086	4 HUWE 64	HBC	$K^- p$ 1.15-1.30 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1383 $\pm$ 1	4.5k	1 BAUBILLIER 79B	HBC	$K^- p$ 8.25 GeV/c
1380 $\pm$ 6	150	1 SUGAHARA 79B	HBC	$\pi^- p$ 6 GeV/c
1387 $\pm$ 3	12k	1,2 BARREIRO 77B	HBC	$K^- p$ 4.2 GeV/c
1391 $\pm$ 3	193	HOLMGREN 77	HBC	See AGUILAR-BENITEZ 81D
1383 $\pm$ 2		1 BARDADIN-... 75	HBC	$K^- p$ 14.3 GeV/c
1389 $\pm$ 1	3060	3 BERTHON 74	HBC	$K^- p$ 1263-1843 MeV/c
1389 $\pm$ 9	15	LONDON 66	HBC	$K^- p$ 2.24 GeV/c
1391.5 $\pm$ 2.6	120	4 SMITH 65	HBC	$K^- p$ 1.8 GeV/c
1399.8 $\pm$ 2.2	58	4 SMITH 65	HBC	$K^- p$ 1.95 GeV/c
1392.0 $\pm$ 6.2	200	COOPER 64	HBC	$K^- p$ 1.45 GeV/c
1382 $\pm$ 3	93	DAHL 61	DBC	$K^- d$ 0.45 GeV/c
1376.0 $\pm$ 4.4	224	4 ELY 61	HLBC	$K^- p$ 1.11 GeV/c



# Baryon Particle Listings

## $\Sigma(1385)$

### $m_{\Sigma(1385)^-} - m_{\Sigma(1385)^+}$

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
- 2 to +6	95	<sup>7</sup> BORENSTEIN	74	HBC $K^- p$ 2.18 GeV/c
7.2±1.4		<sup>7</sup> HABIBI	73	HBC $K^- p \rightarrow \Lambda \pi \pi$
6.3±2.0		<sup>7</sup> SIEGEL	67	HBC $K^- p$ 2.1 GeV/c
11 ±9		<sup>7</sup> LONDON	66	HBC $K^- p$ 2.24 GeV/c
9 ±6		LONDON	66	HBC $\Lambda 3\pi$ events
2.0±1.5		<sup>7</sup> ARMENTEROS65B	HBC	$K^- p$ 0.9-1.2 GeV/c
7.2±2.1		<sup>7</sup> SMITH	65	HBC $K^- p$ 1.8 GeV/c
17.2±2.0		<sup>7</sup> SMITH	65	HBC $K^- p$ 1.95 GeV/c
17 ±7		<sup>7</sup> COOPER	64	HBC $K^- p$ 1.45 GeV/c
4.3±2.2		<sup>7</sup> HUWE	64	HBC $K^- p$ 1.22 GeV/c
0.0±4.2		<sup>7</sup> ELY	61	HLBC $K^- p$ 1.11 GeV/c

### $m_{\Sigma(1385)^0} - m_{\Sigma(1385)^+}$

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-4 to +4	95	<sup>7</sup> BORENSTEIN	74	HBC $K^- p$ 2.18 GeV/c

### $m_{\Sigma(1385)^-} - m_{\Sigma(1385)^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.0±2.4	<sup>7</sup> THOMAS	73	HBC $\pi^- p \rightarrow \Lambda \pi^- K^+$

### $\Sigma(1385)$ WIDTHS

#### $\Sigma(1385)^+$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>36.0± 0.7 OUR AVERAGE</b>				
40.2± 2.1 <sup>+1.2</sup> <sub>-2.8</sub>		AGAKISHIEV	12	SPEC $pp \rightarrow \Sigma(1385)^+ K^+ n$ , 3.5 GeV
37.2± 2.0	1897	BAUBILLIER	84	HBC $K^- p$ 8.25 GeV/c
35.1± 1.7	5256	AGUILAR-...	81D	HBC $K^- p \rightarrow \Lambda \pi \pi$ 4.2 GeV/c
37.5± 2.0	9361	AGUILAR-...	81D	HBC $K^- p \rightarrow \Lambda 3\pi$ 4.2 GeV/c
35.5± 1.9	6900	CAMERON	78	HBC $K^- p$ 0.96-1.36 GeV/c
34.0± 1.6	6846	<sup>8</sup> BORENSTEIN	74	HBC $K^- p$ 2.18 GeV/c
38.3± 3.2	2300	<sup>9</sup> HABIBI	73	HBC $K^- p \rightarrow \Lambda \pi \pi$
32.5± 6.0	400	AGUILAR-...	72B	HBC $K^- p \rightarrow \Lambda \pi$ 's
36 ± 4	1260	<sup>9</sup> SIEGEL	67	HBC $K^- p$ 2.1 GeV/c
32.0± 4.7	750	<sup>9</sup> ARMENTEROS65B	HBC	$K^- p$ 0.95-1.20 GeV/c
46.5± 6.4	859	<sup>9</sup> HUWE	64	HBC $K^- p$ 1.15-1.30 GeV/c

• • • We do not use the following data for averages, fits, limits, etc. • • •				
40 ± 3	600	BAKER	80	HYBR $\pi^+ p$ 7 GeV/c
37 ± 2	750	BAKER	80	HYBR $K^- p$ 7 GeV/c
37 ± 2	7k	<sup>1</sup> BAUBILLIER	79B	HBC $K^- p$ 8.25 GeV/c
30 ± 4	2k	CAUTIS	79	HYBR $\pi^+ p/K^- p$ 11.5 GeV
30 ± 6	100	<sup>1</sup> SUGAHARA	79B	HBC $\pi^- p$ 6 GeV/c
43 ± 5	22k	<sup>1,2</sup> BARREIRO	77B	HBC $K^- p \rightarrow \Sigma \pi$ 's 4 GeV/c
34 ± 2	2594	HOLMGREN	77	HBC See AGUILAR-BENITEZ 81D

40.0± 3.2		<sup>1</sup> BARDADIN-...	75	HBC $K^- p$ 14.3 GeV/c
48 ± 3	3740	<sup>3</sup> BERTHON	74	HBC $K^- p$ 1263-1843 MeV/c
33 ± 20	46	<sup>9</sup> AGUILAR-...	70B	HBC $K^- p \rightarrow \Sigma \pi$ 's 4 GeV/c
25 ± 32	62	<sup>9</sup> BIRMINGHAM	66	HBC $K^- p$ 3.5 GeV/c
30.3± 7.5	250	<sup>9</sup> SMITH	65	HBC $K^- p$ 1.8 GeV/c
33.1± 8.3	250	<sup>9</sup> SMITH	65	HBC $K^- p$ 1.95 GeV/c
51 ± 16	170	<sup>9</sup> COOPER	64	HBC $K^- p$ 1.45 GeV/c
48 ± 16	154	<sup>9</sup> ELY	61	HLBC $K^- p$ 1.11 GeV/c

#### $\Sigma(1385)^0$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>36 ± 5 OUR AVERAGE</b>				
34.8± 5.6	5722	AGUILAR-...	81D	HBC $K^- p \rightarrow \Lambda 3\pi$ 4.2 GeV/c
39.3± 10.2	240	<sup>9</sup> THOMAS	73	HBC $\pi^- p \rightarrow \Lambda \pi^0 K^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
53 ± 8	3100	<sup>10</sup> BORENSTEIN	74	HBC $K^- p \rightarrow \Lambda 3\pi$ 2.18 GeV/c
30 ± 9	106	CURTIS	63	OSPK $\pi^- p$ 1.5 GeV/c

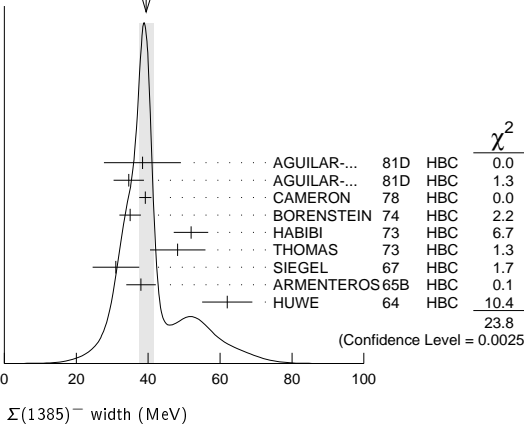
#### $\Sigma(1385)^-$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>39.4± 2.1 OUR AVERAGE</b>				
38.4± 10.7	620	AGUILAR-...	81D	HBC $K^- p \rightarrow \Lambda \pi \pi$ 4.2 GeV/c
34.6± 4.2	3346	AGUILAR-...	81D	HBC $K^- p \rightarrow \Lambda 3\pi$ 4.2 GeV/c
39.2± 1.7	9720	CAMERON	78	HBC $K^- p$ 0.96-1.36 GeV/c
35 ± 3	2303	<sup>8</sup> BORENSTEIN	74	HBC $K^- p$ 2.18 GeV/c
51.9± 4.8	1900	<sup>9</sup> HABIBI	73	HBC $K^- p \rightarrow \Lambda \pi \pi$
48.2± 7.7	630	<sup>9</sup> THOMAS	73	HBC $\pi^- p \rightarrow \Lambda \pi^- K^0$
31.0± 6.5	370	<sup>9</sup> SIEGEL	67	HBC $K^- p$ 2.1 GeV/c
38.0± 4.1	1382	<sup>9</sup> ARMENTEROS65B	HBC	$K^- p$ 0.95-1.20 GeV/c
62 ± 7	1086	HUWE	64	HBC $K^- p$ 1.15-1.30 GeV/c

• • • We do not use the following data for averages, fits, limits, etc. • • •

44 ± 4	4.5k	<sup>1</sup> BAUBILLIER	79B	HBC $K^- p$ 8.25 GeV/c
58 ± 4	150	<sup>1</sup> SUGAHARA	79B	HBC $\pi^- p$ 6 GeV/c
45 ± 5	12k	<sup>1,2</sup> BARREIRO	77B	HBC $K^- p$ 4.2 GeV/c
35 ± 10	193	HOLMGREN	77	HBC See AGUILAR-BENITEZ 81D
47 ± 6		<sup>1</sup> BARDADIN-...	75	HBC $K^- p$ 14.3 GeV/c
40 ± 3	3060	<sup>3</sup> BERTHON	74	HBC $K^- p$ 1263-1843 MeV/c
29.2± 10.6	120	<sup>9</sup> SMITH	65	HBC $K^- p$ 1.80 GeV/c
17.1± 8.9	58	<sup>9</sup> SMITH	65	HBC $K^- p$ 1.95 GeV/c
88 ± 24	200	<sup>9</sup> COOPER	64	HBC $K^- p$ 1.45 GeV/c
40		DAHL	61	DBC $K^- d$ 0.45 GeV/c
66 ± 18	224	<sup>9</sup> ELY	61	HLBC $K^- p$ 1.11 GeV/c

WEIGHTED AVERAGE  
39.4±2.1 (Error scaled by 1.7)



### $\Sigma(1385)$ POLE POSITIONS

#### $\Sigma(1385)^+$ REAL PART

VALUE	DOCUMENT ID	COMMENT
1379±1	LICHTENBERG74	Extrapolates HABIBI 73

#### $\Sigma(1385)^+$ -IMAGINARY PART

VALUE	DOCUMENT ID	COMMENT
17.5±1.5	LICHTENBERG74	Extrapolates HABIBI 73

#### $\Sigma(1385)^-$ REAL PART

VALUE	DOCUMENT ID	COMMENT
1383±1	LICHTENBERG74	Extrapolates HABIBI 73

#### $\Sigma(1385)^-$ -IMAGINARY PART

VALUE	DOCUMENT ID	COMMENT
22.5±1.5	LICHTENBERG74	Extrapolates HABIBI 73

### $\Sigma(1385)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $\Lambda \pi$	(87.0 ± 1.5 ) %	
$\Gamma_2$ $\Sigma \pi$	(11.7 ± 1.5 ) %	
$\Gamma_3$ $\Lambda \gamma$	( 1.25 <sup>+0.13</sup> <sub>-0.12</sub> ) %	
$\Gamma_4$ $\Sigma^+ \gamma$	( 7.0 ± 1.7 ) × 10 <sup>-3</sup>	
$\Gamma_5$ $\Sigma^- \gamma$	< 2.4 × 10 <sup>-4</sup>	90%
$\Gamma_6$ $N \bar{K}$		

The above branching fractions are our estimates, not fits or averages.

### $\Sigma(1385)$ BRANCHING RATIOS

$\Gamma(\Sigma \pi)/\Gamma(\Lambda \pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_2/\Gamma_1$
<b>0.135±0.011 OUR AVERAGE</b>					
0.20 ± 0.06	DIONISI	78B	HBC	± $K^- p \rightarrow Y^* K \bar{K}$	
0.16 ± 0.03	BERTHON	74	HBC	+ $K^- p$ 1.26-1.84 GeV/c	
0.11 ± 0.02	BERTHON	74	HBC	- $K^- p$ 1.26-1.84 GeV/c	
0.21 ± 0.05	BORENSTEIN	74	HBC	+ $K^- p \rightarrow \Lambda \pi^+ \pi^-$ , $\Sigma^0 \pi^+ \pi^-$	
0.18 ± 0.04	MAST	73	MPWA	± $K^- p \rightarrow \Lambda \pi^+ \pi^-$ , $\Sigma^0 \pi^+ \pi^-$	
0.10 ± 0.05	THOMAS	73	HBC	- $\pi^- p \rightarrow \Lambda K \pi$ , $\Sigma K \pi$	



See key on page 601

Baryon Particle Listings  
 $\Sigma(1385)$ ,  $\Sigma(1480)$  Bumps

0.16 ± 0.07	AGUILAR-...	72B	HBC	+	$K^- p$ 3.9, 4.6 GeV/c
0.13 ± 0.04	COLLEY	71B	DBC	→ 0	$K^- N$ 1.5 GeV/c
0.13 ± 0.04	PAN	69	HBC	+	$\pi^+ p \rightarrow \Lambda K \pi$ , $\Sigma K \pi$
0.08 ± 0.06	LONDON	66	HBC	+	$K^- p$ 2.24 GeV/c
0.163 ± 0.041	ARMENTEROS65B	HBC	±		$K^- p$ 0.95–1.20 GeV/c
0.09 ± 0.04	HUWE	64	HBC	±	$K^- p$ 1.2–1.7 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.04	ALSTON	62	HBC	± 0	$K^- p$ 1.15 GeV/c
0.04 ± 0.04	BASTIEN	61	HBC	±	

$\Gamma(\Lambda\gamma)/\Gamma(\Lambda\pi)$   $\Gamma_3/\Gamma_1$

This ratio is of course for $\Sigma(1385)^0 \rightarrow \Lambda\gamma$ and $\Lambda\pi^0$ .					
VALUE (units $10^{-2}$ )	EVTs	DOCUMENT ID	TECN	COMMENT	
<b><math>1.43^{+0.15}_{-0.13}</math></b>	<b>OUR AVERAGE</b>				
$1.42 \pm 0.12^{+0.11}_{-0.07}$	$624 \pm 25$	KELLER	11	CLAS	$\gamma p \rightarrow K^+ \Lambda\gamma$ , $E_\gamma$ 1.6–3.8 GeV
$1.53 \pm 0.39^{+0.15}_{-0.24}$	61	TAYLOR	05	CLAS	$\gamma p \rightarrow K^+ \Lambda\gamma$

$\Gamma(\Sigma^+\gamma)/\Gamma(\Sigma\pi)$   $\Gamma_4/\Gamma_2$

This ratio is for $\Sigma(1385)^+ \rightarrow \Sigma^+\gamma$ over $\Sigma(1385)^+ \rightarrow \Sigma\pi$ .					
VALUE (%)	DOCUMENT ID	TECN	COMMENT		
<b>5.98 ± 1.11<sup>+0.27</sup><sub>-0.61</sub></b>	11	KELLER	12	CLAS	$\gamma p \rightarrow K^0 \Sigma(1385)^+$

$\Gamma(\Sigma^-\gamma)/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	CHG.	COMMENT
<b>&lt;2.4 × 10<sup>-4</sup></b>	90	12	MOLCHANOV	04	SELX — $\Sigma^- \text{Pb} \rightarrow \Sigma(1385)^- \text{Pb}$ , 600 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<6.1 × 10 <sup>-4</sup>	90	13	ARIK	77	SPEC — $\Sigma^- \text{Pb} \rightarrow \Sigma(1385)^- \text{Pb}$ , 23 GeV

$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Sigma(1385) \rightarrow \Lambda\pi$   $(\Gamma_6\Gamma_1)^{1/2}/\Gamma$

VALUE	DOCUMENT ID	CHG.	COMMENT
+0.586 ± 0.319	14	DEVENISH	74B 0 Fixed-t dispersion rel.

$\Sigma(1385)$  FOOTNOTES

- <sup>1</sup> From fit to inclusive  $\Lambda\pi$  spectrum.  
<sup>2</sup> Includes data of HOLMGREN 77.  
<sup>3</sup> The errors are statistical only. The resolution is not unfolded.  
<sup>4</sup> The error is enlarged to  $\Gamma/\sqrt{N}$ . See the note on the  $K^*(892)$  mass in the 1984 edition.  
<sup>5</sup> From a fit to  $\Lambda\pi^0$  with the width fixed at 34 MeV.  
<sup>6</sup> From fit to inclusive  $\Lambda\pi^0$  spectrum with the width fixed at 40 MeV.  
<sup>7</sup> Redundant with data in the mass Listings.  
<sup>8</sup> Results from  $\Lambda\pi^+\pi^-$  and  $\Lambda\pi^+\pi^-\pi^0$  combined by us.  
<sup>9</sup> The error is enlarged to  $4\Gamma/\sqrt{N}$ . See the note on the  $K^*(892)$  mass in the 1984 edition.  
<sup>10</sup> Consistent with +, 0, and – widths equal.  
<sup>11</sup> KELLER 12 gives  $\Gamma(\Sigma^+\gamma)/\Gamma(\Sigma^+\pi^0) = (11.95 \pm 2.21^{+0.53}_{-1.21})\%$ , using 1/2 our total  $\Sigma(1385) \rightarrow \Sigma\pi$  fraction for  $\Sigma^+\pi^0$ . We divide the KELLER 12 value by two.  
<sup>12</sup> We calculate this from the MOLCHANOV 04 upper limit of 9.5 keV on the  $\Sigma^-\gamma$  width.  
<sup>13</sup> We calculate this from the ARIK 77 upper limit of 24 keV on the  $\Sigma^-\gamma$  width.  
<sup>14</sup> An extrapolation of the parametrized amplitude below threshold.

$\Sigma(1385)$  REFERENCES

AGAKISHIEV	12	PR C95 035203	G. Agakishiev et al.	(HADES Collab.)
KELLER	12	PR D85 052004	D. Keller et al.	(JLab CLAS Collab.)
KELLER	11	PR D83 072004	D. Keller et al.	(JLab CLAS Collab.)
TAYLOR	05	PR C71 054609	S. Taylor et al.	(JLab CLAS Collab.)
Also		PR C72 039902 (errat.)	S. Taylor et al.	(JLab CLAS Collab.)
MOLCHANOV	04	PL B590 161	V.V. Molchanov et al.	(FNAL SELEX Collab.)
BAUBILLIER	84	ZPHY C23 213	M. Baubillier et al.	(BIRM, CERN, GLAS+)
PDG	84	RMP 56 51	C.G. Wohl et al.	(LBL, CIT, CERN)
AGUILAR-...	81D	AFIS A77 144	M. Aguilar-Benitez, J. Salicio	(MADR)
BAKER	80	NP B166 207	P.A. Baker et al.	(LOIC)
BAUBILLIER	79B	NP B148 18	M. Baubillier et al.	(BIRM, CERN, GLAS+)
CAUTIS	79	NP B156 507	C.V. Cautis et al.	(SLAC)
SUGAHARA	79B	NP B156 237	R. Sugahara et al.	(KEK, OSKC, KINK)
CAMERON	78	NP B143 189	W. Cameron et al.	(RHEL, LOIC)
DIONISI	78B	PL 78B 154	C. Dionisi, R. Armenteros, J. Diaz	(CERN, AMST+)
ARIK	77	PRL 38 1000	E. Arik et al.	(PITT, BNL, MASA)
BARREIRO	77B	NP B126 319	F. Barreiro et al.	(CERN, AMST, NUM)
HOLMGREN	77	NP B119 261	S.O. Holmgren et al.	(CERN, AMST, NUM)
BARDADIN-...	75	NP B98 418	M. Bardadin-Otwinowska et al.	(SACL, EPOL+)
BERTHON	74	NC 21A 146	A. Berthon et al.	(CDEF, RHEL, SACL+)
BORNSTEIN	74	PR D9 3006	S.R. Bornstein et al.	(BNL, MICH)
DEVENISH	74B	NP B81 330	R.C.E. Devenish, C.D. Froggatt, B.R. Martin	(DESY+)
LICHTENBERG	74	PR D10 3865	D.B. Lichtenberg	(IND)
Also		Private Comm.	M. Habibi	(COLU)
HABIBI	73	Thesis Nevis 199	C. Baltay et al.	(COLU, BING)
Also		Purdue Conf. 387	T.S. Mast et al.	(LBL) IUP
MAST	73	PR D7 3212	T.S. Mast et al.	(LBL) IUP
Also		PR D7 5	D.W. Thomas et al.	(CMU) JP
THOMAS	73	NP B56 15	M. Aguilar-Benitez et al.	(BNL)
AGUILAR-...	72B	PR D6 29	D.C. Colley et al.	(BIRM, EDIN, GLAS+)
COLLEY	71B	NP B31 61	M. Aguilar-Benitez et al.	(BNL, SYRA)
AGUILAR-...	70B	PRL 25 58	Y.L. Pan, F.L. Forman	(PERN) I
PAN	69	PRL 23 808	D.M. Siegel	(LRL)
SIEGEL	66	Thesis UCRL 18041	M. Haque et al.	(BIRM, GLAS, LOIC, OXF+)
BIRMINGHAM	66	PR 152 1148	G.W. London et al.	(BNL, SYRA) J
LONDON	66	PR 143 1034		

ARMENTEROS	65B	PL 19 75	R. Armenteros et al.	(CERN, HEID, SACL)
SMITH	65	Thesis UCLA	L.T. Smith	(UCLA)
COOPER	64	PL 8 365	W.A. Cooper et al.	(CERN, AMST)
HUWE	64	Thesis UCRL 11291	D.O. Huwe	(LRL) JP
Also		PR 181 1824	D.O. Huwe	(LRL)
CURTIS	63	PR 132 1771	L.J. Curtis et al.	(MICH) J
ALSTON	62	CERN Conf. 311	M.H. Alston et al.	(LRL)
BASTIEN	61	PRL 6 702	P.L. Bastien, M. Ferro-Luzzi, A.H. Rosenfeld	(LRL)
DAHL	61	PRL 6 142	O.I. Dahl et al.	(LRL)
ELY	61	PRL 7 461	R.P. Ely et al.	(LRL) J
ALSTON	60	PRL 5 520	M.H. Alston et al.	(LRL) I

$\Sigma(1480)$  Bumps  $I(J^P) = 1(?)^?$  Status: \*

OMITTED FROM SUMMARY TABLE  
These are peaks seen in  $\Lambda\pi$  and  $\Sigma\pi$  spectra in the reaction  $\pi^+ p \rightarrow (\gamma\pi) K^+$  at 1.7 GeV/c. Also, the  $\gamma$  polarization oscillates in the same region.

MILLER 70 suggests a possible alternate explanation in terms of a reflection of  $N(1675) \rightarrow \Lambda K$  decay. However, such an explanation for the  $(\Sigma^+\pi^0) K^+$  channel in terms of  $\Delta(1650) \rightarrow \Sigma K$  decay seems unlikely (see PAN 70). In addition such reflections would also have to account for the oscillation of the  $\gamma$  polarization in the 1480 MeV region.

HANSON 71, with less data than PAN 70, can neither confirm nor deny the existence of this state. MAST 75 sees no structure in this region in  $K^- p \rightarrow \Lambda\pi^0$ .

ENGELN 80 performs a multichannel analysis of  $K^- p \rightarrow p\bar{K}^0\pi^-$  at 4.2 GeV/c. They observe a 3.5 standard-deviation signal at 1480 MeV in  $p\bar{K}^0$  which cannot be explained as a reflection of any competing channel.

PRAKHOV 04 sees no evidence for this or other light  $\Sigma$  resonances, aside from the  $\Sigma(1385)$ , in  $K^- p \rightarrow \Lambda\pi^0\pi^0$ .

ZYCHOR 06 finds peaks in  $pp \rightarrow pK^+(\pi^\pm X^\mp)$  at  $p_{\text{beam}} = 3.65$  GeV/c.

$\Sigma(1480)$  MASS  
(PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>≈ 1480 OUR ESTIMATE</b>				
1480 ± 15	365 ± 60	ZYCHOR	06	SPEC $pp \rightarrow pK^+(\pi^\pm X^\mp)$
1480	120	ENGELN	80	HBC $K^- p \rightarrow (p\bar{K}^0)\pi^-$
1485 ± 10		CLINE	73	MPWA $K^- d \rightarrow (\Lambda\pi^-)p$
1479 ± 10		PAN	70	HBC $\pi^+ p \rightarrow (\Lambda\pi^+)K^+$
1465 ± 15		PAN	70	HBC $\pi^+ p \rightarrow (\Sigma\pi)K^+$

$\Sigma(1480)$  WIDTH  
(PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
60 ± 15	365 ± 60	ZYCHOR	06	SPEC $pp \rightarrow pK^+(\pi^\pm X^\mp)$
80 ± 20	120	ENGELN	80	HBC $K^- p \rightarrow (p\bar{K}^0)\pi^-$
40 ± 20		CLINE	73	MPWA $K^- d \rightarrow (\Lambda\pi^-)p$
31 ± 15		PAN	70	HBC $\pi^+ p \rightarrow (\Lambda\pi^+)K^+$
30 ± 20		PAN	70	HBC $\pi^+ p \rightarrow (\Sigma\pi)K^+$

$\Sigma(1480)$  DECAY MODES  
(PRODUCTION EXPERIMENTS)

Mode	
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Lambda\pi$
$\Gamma_3$	$\Sigma\pi$

$\Sigma(1480)$  BRANCHING RATIOS  
(PRODUCTION EXPERIMENTS)

$\Gamma(\Sigma\pi)/\Gamma(\Lambda\pi)$ $\Gamma_3/\Gamma_2$			
VALUE	DOCUMENT ID	TECN	CHG.
0.82 ± 0.51	PAN	70	HBC +
$\Gamma(N\bar{K})/\Gamma(\Lambda\pi)$ $\Gamma_1/\Gamma_2$			
VALUE	DOCUMENT ID	TECN	CHG.
0.72 ± 0.50	PAN	70	HBC +
$\Gamma(N\bar{K})/\Gamma_{\text{total}}$ $\Gamma_1/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT
small	CLINE	73	MPWA $K^- d \rightarrow (\Lambda\pi^-)p$

Baryon Particle Listings

$\Sigma(1480)$  Bumps,  $\Sigma(1560)$  Bumps,  $\Sigma(1580)$ ,  $\Sigma(1620)$

$\Sigma(1480)$ REFERENCES (PRODUCTION EXPERIMENTS)					
ZYCHOR	06	PRL 96 012002	I. Zychor <i>et al.</i>	(ANKE Collab.)	
PRAKHOV	04	PR C69 042202	S. Prakhov <i>et al.</i>	(BNL Crystal Ball Collab.)	
ENGELÉN	80	NP B167 61	J.J. Engelen <i>et al.</i>	(NUM, AMST, CERN+)	
MAST	75	PR D11 3078	T.S. Mast <i>et al.</i>	(LBL)	
CLINE	73	LNC 6 205	D. Cline, R. Laumann, J. Mapp	(WISC) IJP	
HANSON	71	PR D4 1296	P. Hanson, G.E. Kalmus, J. Louie	(LBL) I	
MILLER	70	Duke Conf. 229	D.H. Miller	(PURD)	
Hyperon Resonances, 1970					
PAN	70	PR D2 449	Y.L. Pan <i>et al.</i>	(PENN)	
Also		PR L 23 808	Y.L. Pan, F.L. Forman	(PENN) I	
Also		PR L 23 806	Y.L. Pan, F.L. Forman	(PENN) I	

$\Sigma(1560)$ Bumps	$I(J^P) = 1(?^?)$	Status: **
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OMITTED FROM SUMMARY TABLE

This entry lists peaks reported in mass spectra around 1560 MeV without implying that they are necessarily related.

DIONISI 78B observes a 6 standard-deviation enhancement at 1553 MeV in the charged  $\Lambda/\Sigma\pi$  mass spectra from  $K^-p \rightarrow (\Lambda/\Sigma)\pi K\bar{K}$  at 4.2 GeV/c. In a CERN ISR experiment, LOCKMAN 78 reports a narrow 6 standard-deviation enhancement at 1572 MeV in  $\Lambda\pi^\pm$  from the reaction  $pp \rightarrow \Lambda\pi^+\pi^-X$ . These enhancements are unlikely to be associated with the  $\Sigma(1580)$  (which has not been confirmed by several recent experiments – see the next entry in the Listings).

CARROLL 76 observes a bump at 1550 MeV (as well as one at 1580 MeV) in the isospin-1  $\bar{K}N$  total cross section, but uncertainties in cross section measurements outside the mass range of the experiment preclude estimating its significance.

See also MEADOWS 80 for a review of this state.

$\Sigma(1560)$ MASS (PRODUCTION EXPERIMENTS)					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$\approx 1560$ OUR ESTIMATE					
1553 $\pm$ 7	121	DIONISI	78B	HBC	$\pm$ $K^-p \rightarrow (Y\pi)K\bar{K}$
1572 $\pm$ 4	40	LOCKMAN	78	SPEC	$\pm$ $pp \rightarrow \Lambda\pi^+\pi^-X$

$\Sigma(1560)$ WIDTH (PRODUCTION EXPERIMENTS)					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
79 $\pm$ 30	121	DIONISI	78B	HBC	$\pm$ $K^-p \rightarrow (Y\pi)K\bar{K}$
15 $\pm$ 6	40	<sup>1</sup> LOCKMAN	78	SPEC	$\pm$ $pp \rightarrow \Lambda\pi^+\pi^-X$

$\Sigma(1560)$ DECAY MODES (PRODUCTION EXPERIMENTS)	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\Lambda\pi$	seen
$\Gamma_2$ $\Sigma\pi$	

$\Sigma(1560)$ BRANCHING RATIOS (PRODUCTION EXPERIMENTS)					
$\Gamma(\Sigma\pi)/[\Gamma(\Lambda\pi) + \Gamma(\Sigma\pi)]$					$\Gamma_2/(\Gamma_1 + \Gamma_2)$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
0.35 $\pm$ 0.12	DIONISI	78B	HBC	$\pm$ $K^-p \rightarrow (Y\pi)K\bar{K}$	
$\Gamma(\Lambda\pi)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
seen	LOCKMAN	78	SPEC	$\pm$ $pp \rightarrow \Lambda\pi^+\pi^-X$	

$\Sigma(1560)$ FOOTNOTES (PRODUCTION EXPERIMENTS)					
<sup>1</sup> The width observed by LOCKMAN 78 is consistent with experimental resolution.					
$\Sigma(1560)$ REFERENCES (PRODUCTION EXPERIMENTS)					
MEADOWS	80	Toronto Conf. 283	B.T. Meadows	(CINIC)	
DIONISI	78B	PL 78B 154	C. Dionisi, R. Armenteros, J. Diaz	(CERN, AMST+)	I
LOCKMAN	78	Saclay DPHFE 78-01	W. Lockman <i>et al.</i>	(UCLA, SACL)	
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL)	I

$\Sigma(1580) 3/2^-$	$I(J^P) = 1(\frac{3}{2}^-)$	Status: *
OMITTED FROM SUMMARY TABLE		
Seen in the isospin-1 $\bar{K}N$ cross section at BNL (LI 73, CARROLL 76) and in a partial-wave analysis of $K^-p \rightarrow \Lambda\pi^0$ for c.m. energies 1560–1600 MeV by LITCHFIELD 74. LITCHFIELD 74 finds $J^P = 3/2^-$ . Not seen by ENGLER 78 or by CAMERON 78C (with larger statistics in $K_L^0 p \rightarrow \Lambda\pi^+$ and $\Sigma^0\pi^+$ ).		
Neither OLMSTED 04 (in $K^-p \rightarrow \Lambda\pi^0$ ) nor PRAKHOV 04 (in $K^-p \rightarrow \Lambda\pi^0\pi^0$ ) see any evidence for this state.		

$\Sigma(1580)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 1580$ OUR ESTIMATE			
1583 $\pm$ 4	<sup>1</sup> CARROLL	76	DPWA Isospin-1 total $\sigma$
1582 $\pm$ 4	<sup>2</sup> LITCHFIELD	74	DPWA $K^-p \rightarrow \Lambda\pi^0$

$\Sigma(1580)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
15	<sup>1</sup> CARROLL	76	DPWA Isospin-1 total $\sigma$
11 $\pm$ 4	<sup>2</sup> LITCHFIELD	74	DPWA $K^-p \rightarrow \Lambda\pi^0$

$\Sigma(1580)$ DECAY MODES	
Mode	
$\Gamma_1$ $N\bar{K}$	
$\Gamma_2$ $\Lambda\pi$	
$\Gamma_3$ $\Sigma\pi$	

$\Sigma(1580)$ BRANCHING RATIOS	
See “Sign conventions for resonance couplings” in the Note on $\Lambda$ and $\Sigma$ Resonances.	

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
+0.03 $\pm$ 0.01	<sup>2</sup> LITCHFIELD	74	DPWA	$\bar{K}N$ multichannel	
$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1580) \rightarrow \Lambda\pi$					$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
not seen	CAMERON	78C	HBC	$K_L^0 p \rightarrow \Lambda\pi^+$	
not seen	ENGLER	78	HBC	$K_L^0 p \rightarrow \Lambda\pi^+$	
+0.10 $\pm$ 0.02	<sup>2</sup> LITCHFIELD	74	DPWA	$K^-p \rightarrow \Lambda\pi^0$	

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1580) \rightarrow \Sigma\pi$					$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
not seen	CAMERON	78C	HBC	$K_L^0 p \rightarrow \Sigma^0\pi^+$	
not seen	ENGLER	78	HBC	$K_L^0 p \rightarrow \Sigma^0\pi^+$	
+0.03 $\pm$ 0.04	<sup>2</sup> LITCHFIELD	74	DPWA	$\bar{K}N$ multichannel	

$\Sigma(1580)$ FOOTNOTES	
<sup>1</sup> CARROLL 76 sees a total-cross-section bump with $(J+1/2)\Gamma_{\text{el}}/\Gamma_{\text{total}} = 0.06$ .	
<sup>2</sup> The main effect observed by LITCHFIELD 74 is in the $\Lambda\pi$ final state; the $\bar{K}N$ and $\Sigma\pi$ couplings are estimated from a multichannel fit including total-cross-section data of LI 73.	

$\Sigma(1580)$ REFERENCES					
OLMSTED	04	PL B588 29	J. Olmsted <i>et al.</i>	(BNL Crystal Ball Collab.)	
PRAKHOV	04	PR C69 042202	S. Prakhov <i>et al.</i>	(BNL Crystal Ball Collab.)	
CAMERON	78C	NP B132 189	W. Cameron <i>et al.</i>	(BGNA, EDIN, GLAS+)	I
ENGLER	78	PR D18 3061	A. Engler <i>et al.</i>	(CMU, ANL)	
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL)	I
LITCHFIELD	74	PL 51B 509	P.J. Litchfield	(CERN) IJP	
LI	73	Purdue Conf. 283	K.K. Li	(BNL)	I

$\Sigma(1620) 1/2^-$	$I(J^P) = 1(\frac{1}{2}^-)$	Status: *
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OMITTED FROM SUMMARY TABLE

The  $S_{11}$  state at 1697 MeV reported by VANHORN 75 is tentatively listed under the  $\Sigma(1750)$ . CARROLL 76 sees two bumps in the isospin-1 total cross section near this mass. GAO 12 sees no evidence for this resonance.

Production experiments are listed separately in the next entry.

See key on page 601

Baryon Particle Listings

$\Sigma(1620)$ ,  $\Sigma(1620)$  Production Experiments

$\Sigma(1620)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 1620$ OUR ESTIMATE			
$1600 \pm 15$	ZHANG	13A	DPWA Multichannel
$1600 \pm 6$	<sup>1</sup> MORRIS	78	DPWA $K^- n \rightarrow \Lambda \pi^-$
$1608 \pm 5$	<sup>2</sup> CARROLL	76	DPWA Isospin-1 total $\sigma$
$1633 \pm 10$	<sup>3</sup> CARROLL	76	DPWA Isospin-1 total $\sigma$
$1630 \pm 10$	LANGBEIN	72	IPWA $\overline{K} N$ multichannel
1620	KIM	71	DPWA K-matrix analysis

$\Sigma(1620)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$400 \pm 152$	ZHANG	13A	DPWA Multichannel
$87 \pm 19$	<sup>1</sup> MORRIS	78	DPWA $K^- n \rightarrow \Lambda \pi^-$
15	<sup>2</sup> CARROLL	76	DPWA Isospin-1 total $\sigma$
10	<sup>3</sup> CARROLL	76	DPWA Isospin-1 total $\sigma$
$65 \pm 20$	LANGBEIN	72	IPWA $\overline{K} N$ multichannel
40	KIM	71	DPWA K-matrix analysis

$\Sigma(1620)$  POLE POSITION

REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1501	ZHANG	13A	DPWA Multichannel

−2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
171	ZHANG	13A	DPWA Multichannel

$\Sigma(1620)$  DECAY MODES

Mode
$\Gamma_1 \quad N \overline{K}$
$\Gamma_2 \quad \Lambda \pi$
$\Gamma_3 \quad \Sigma \pi$

$\Sigma(1620)$  BRANCHING RATIOS

$\Gamma(N \overline{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE				
$0.59 \pm 0.10$	ZHANG	13A	DPWA Multichannel	
$0.22 \pm 0.02$	LANGBEIN	72	IPWA $\overline{K} N$ multichannel	
0.05	KIM	71	DPWA K-matrix analysis	

$(\Gamma_1 \Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N \overline{K} \rightarrow \Sigma(1620) \rightarrow \Lambda \pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1 \Gamma_2)^{1/2}/\Gamma$
VALUE				
$0.12 \pm 0.02$	<sup>1</sup> MORRIS	78	DPWA $K^- n \rightarrow \Lambda \pi^-$	
not seen	BAILLON	75	IPWA $\overline{K} N \rightarrow \Lambda \pi$	
0.15	KIM	71	DPWA K-matrix analysis	

$(\Gamma_1 \Gamma_3)^{1/2}/\Gamma_{\text{total}}$ in $N \overline{K} \rightarrow \Sigma(1620) \rightarrow \Sigma \pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1 \Gamma_3)^{1/2}/\Gamma$
VALUE				
$+0.32 \pm 0.03$	ZHANG	13A	DPWA Multichannel	
not seen	HEPP	76B	DPWA $K^- N \rightarrow \Sigma \pi$	
$+0.40 \pm 0.06$	LANGBEIN	72	IPWA $\overline{K} N$ multichannel	
$+0.08$	KIM	71	DPWA K-matrix analysis	

$\Sigma(1620)$  FOOTNOTES

- <sup>1</sup> MORRIS 78 obtains an equally good fit without including this resonance.  
<sup>2</sup> Total cross-section bump with  $(J+1/2) \Gamma_{\text{el}} / \Gamma_{\text{total}}$  is 0.06 seen by CARROLL 76.  
<sup>3</sup> Total cross-section bump with  $(J+1/2) \Gamma_{\text{el}} / \Gamma_{\text{total}}$  is 0.04 seen by CARROLL 76.

$\Sigma(1620)$  REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GAO	12	PR C86 025201	P. Gao, J. Shi, B.S. Zou	(BHEP, BEIJT)
Also		NP A867 41	P. Gao, B.S. Zou, A. Sibirtsev	(BHEP, BEIJT+)
MORRIS	78	PR D17 55	W.A. Morris <i>et al.</i>	(FSU) IJP
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL) I
HEPP	76B	PL 65B 487	V. Hepp <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
VANHORN	75	NP B87 145	A.J. van Horn	(LBL) IJP
Also		NP B87 157	A.J. van Horn	(LBL) IJP
LANGBEIN	72	NP B47 477	W. Langbein, F. Wagner	(MPIM) IJP
KIM	71	PRL 27 356	J.K. Kim	(HARV) IJP
Also		Duke Conf. 161	J.K. Kim	(HARV) IJP
Hyperon Resonances,		1970		

$\Sigma(1620)$  Production Experiments

$I(J^P) = 1(?^?)$

OMITTED FROM SUMMARY TABLE

Formation experiments are listed separately in the previous entry.

The results of CRENNELL 69B at 3.9 GeV/c are not confirmed by SABRE 70 at 3.0 GeV/c. However, at 4.5 GeV/c, AMMANN 70 sees a peak at 1642 MeV which on the basis of branching ratios they do not associate with the  $\Sigma(1670)$ . See MILLER 70 for a review of these conflicts.

$\Sigma(1620)$  MASS  
(PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$\approx 1620$ OUR ESTIMATE					
$1642 \pm 12$		AMMANN	70	DBC	$K^- N$ 4.5 GeV/c
$1618 \pm 3$	20	BLUMENFELD	69	HBC +	$K_L^0 p$
$1619 \pm 8$		CRENNELL	69B	DBC ±	$K^- N \rightarrow \Lambda \pi \pi \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1616 \pm 8$		CRENNELL	68	DBC ±	See CRENNELL 69B

$\Sigma(1620)$  WIDTH  
(PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$55 \pm 24$		AMMANN	70	DBC	$K^- N$ 4.5 GeV/c
$30 \pm 10$	20	BLUMENFELD	69	HBC +	
$72 \pm^{22}_{15}$		CRENNELL	69B	DBC ±	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$66 \pm 16$		CRENNELL	68	DBC ±	See CRENNELL 69B

$\Sigma(1620)$  DECAY MODES  
(PRODUCTION EXPERIMENTS)

Mode
$\Gamma_1 \quad N \overline{K}$
$\Gamma_2 \quad \Lambda \pi$
$\Gamma_3 \quad \Sigma \pi$
$\Gamma_4 \quad \Lambda \pi \pi$
$\Gamma_5 \quad \Sigma(1385) \pi$
$\Gamma_6 \quad \Lambda(1405) \pi$

$\Sigma(1620)$  BRANCHING RATIOS  
(PRODUCTION EXPERIMENTS)

$\Gamma(\Lambda \pi \pi)/\Gamma(\Lambda \pi)$	DOCUMENT ID	TECN	CHG	$\Gamma_4/\Gamma_2$
VALUE				
$\sim 2.5$	14	BLUMENFELD	69	HBC +

$\Gamma(N \overline{K})/\Gamma(\Lambda \pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_1/\Gamma_2$
VALUE					
$0.4 \pm 0.4$	AMMANN	70	DBC	$K^- p$ 4.5 GeV/c	
$0.0 \pm 0.1$	CRENNELL	68	DBC +	See CRENNELL 69B	

$\Gamma(\Lambda \pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	CHG	$\Gamma_2/\Gamma$
VALUE				
large	CRENNELL	68	DBC ±	

$\Gamma(\Sigma(1385) \pi)/\Gamma(\Lambda \pi)$	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_5/\Gamma_2$
VALUE					
$<0.3$	95	AMMANN	70	DBC	$K^- p$ 4.5 GeV/c
$0.2 \pm 0.1$		CRENNELL	68	DBC ±	

$\Gamma(\Sigma\pi)/\Gamma(\Lambda\pi)$					$\Gamma_3/\Gamma_2$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<1.1	95	AMMANN	70	DBC	$K^- N$ 4.5 GeV/c

$\Gamma(\Lambda(1405)\pi)/\Gamma(\Lambda\pi)$					$\Gamma_6/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT		
$0.7 \pm 0.4$	AMMANN	70	DBC	$K^- p$ 4.5 GeV/c	

# Baryon Particle Listings

## $\Sigma(1620)$ Production Experiments, $\Sigma(1660)$ , $\Sigma(1670)$

$\Sigma(1620)$ REFERENCES (PRODUCTION EXPERIMENTS)				
AMMANN	70	PRL 24 327	A.C. Ammann <i>et al.</i>	(PURD, IND)
Also		PR D7 1345	A.C. Ammann <i>et al.</i>	(PURD, IUPU)
MILLER	70	Duke Conf. 229	D.H. Miller	(PURD)
Hyperon Resonances, 1970				
SABRE	70	NP B16 201	R. Barloutaud <i>et al.</i>	(SABRE Collab.)
BLUMENFELD	69	PL 29B 58	B.J. Blumenfeld, G.R. Kalbfleisch	(BNL)1
CRENNELL	69B	Lund Paper 183	D.J. Crennell <i>et al.</i>	(BNL, CUNY)1
Results are quoted in LEVI-SETTI 69C.				
Also		Lund Conf.	R. Levi-Setti	(EFI)
CRENNELL	68	PRL 21 648	D.J. Crennell <i>et al.</i>	(BNL, CUNY)1

$\Sigma(1660) 1/2^+$	$I(J^P) = 1(\frac{1}{2}^+)$ Status: ***
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For results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

$\Sigma(1660)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1630 to 1690 (<math>\approx 1660</math>) OUR ESTIMATE</b>				
1633 $\pm$ 3	GAO	12	DPWA	$\overline{K}N \rightarrow \Lambda\pi$
1665.1 $\pm$ 11.2	<sup>1</sup> KOISO	85	DPWA	$K^-p \rightarrow \Sigma\pi$
1670 $\pm$ 10	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1679 $\pm$ 10	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1676 $\pm$ 15	GOPAL	77	DPWA	$\overline{K}N$ multichannel
1668 $\pm$ 25	VANHORN	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
1670 $\pm$ 20	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1565 or 1597	<sup>2</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
1660 $\pm$ 30	<sup>3</sup> BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$
1671 $\pm$ 2	<sup>4</sup> PONTE	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$

$\Sigma(1660)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>40 to 200 (<math>\approx 100</math>) OUR ESTIMATE</b>				
121 $\pm$ 4	GAO	12	DPWA	$\overline{K}N \rightarrow \Lambda\pi$
— 7				
81.5 $\pm$ 22.2	<sup>1</sup> KOISO	85	DPWA	$K^-p \rightarrow \Sigma\pi$
152 $\pm$ 20	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
38 $\pm$ 10	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
120 $\pm$ 20	GOPAL	77	DPWA	$\overline{K}N$ multichannel
230 $\pm$ 165	VANHORN	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
— 60				
250 $\pm$ 110	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
202 or 217	<sup>2</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
80 $\pm$ 40	<sup>3</sup> BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$
81 $\pm$ 10	<sup>4</sup> PONTE	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$

$\Sigma(1660)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\overline{K}$	10–30 %	
$\Gamma_2$ $\Lambda\pi$	seen	
$\Gamma_3$ $\Sigma\pi$	seen	

$\Sigma(1660)$ BRANCHING RATIOS	
See “Sign conventions for resonance couplings” in the Note on $\Lambda$ and $\Sigma$ Resonances.	

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.1 to 0.3 OUR ESTIMATE</b>			
0.12 $\pm$ 0.03	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$
0.10 $\pm$ 0.05	ALSTON-...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.04	GOPAL	77	DPWA See GOPAL 80
0.27 or 0.29	<sup>2</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1660) \rightarrow \Lambda\pi$		$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
— 0.064 $\pm$ 0.005	GAO	12	DPWA $\overline{K}N \rightarrow \Lambda\pi$
— 0.003			
< 0.04	GOPAL	77	DPWA $\overline{K}N$ multichannel
0.12 $\pm$ 0.12	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
— 0.04			
• • • We do not use the following data for averages, fits, limits, etc. • • •			

— 0.10 or — 0.11	<sup>2</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
— 0.04 $\pm$ 0.02	<sup>3</sup> BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$
+ 0.16 $\pm$ 0.01	<sup>4</sup> PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$

$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1660) \rightarrow \Sigma\pi$		$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
— 0.13 $\pm$ 0.04	<sup>1</sup> KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$
— 0.16 $\pm$ 0.03	GOPAL	77	DPWA $\overline{K}N$ multichannel
— 0.11 $\pm$ 0.01	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
— 0.34 or — 0.37	<sup>2</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
not seen	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$

$\Sigma(1660)$ FOOTNOTES	
<sup>1</sup> The evidence of KOISO 85 is weak.	
<sup>2</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.	
<sup>3</sup> From solution 1 of BAILLON 75; not present in solution 2.	
<sup>4</sup> From solution 2 of PONTE 75; not present in solution 1.	

$\Sigma(1660)$ REFERENCES				
GAO	12	PR C86 025201	P. Gao, J. Shi, B.S. Zou	(BHEP, BEUT)
Also		NP A867 41	P. Gao, B.S. Zou, A. Sibirtsev	(BHEP, BEUT+)
KOISO	85	NP A433 619	H. Koiso <i>et al.</i>	(TOKY, MASA)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HELSE, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IUP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IUP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IUP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IUP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IUP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC) IUP
Also		NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IUP
HEPP	76B	PL 65B 487	V. Hepp <i>et al.</i>	(CERN, HEIDH, MPIM) IUP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IUP
PONTE	75	PR D12 2597	R.A. Ponte <i>et al.</i>	(MASA, TENN, UCR) IUP
VANHORN	75	NP B87 145	A.J. van Horn	(LBL) IUP
Also		NP B87 157	A.J. van Horn	(LBL) IUP
KANE	74	LBL-2452	D.F. Kane	(LBL) IUP

## THE $\Sigma(1670)$ REGION

**Production experiments:** The measured  $\Sigma\pi/\Sigma\pi\pi$  branching ratio for the  $\Sigma(1670)$  produced in the reaction  $K^-p \rightarrow \pi^-\Sigma(1670)^+$  is strongly dependent on momentum transfer. This was first discovered by EBERHARD 69, who suggested that there exist two  $\Sigma$  resonances with the same mass and quantum numbers: one with a large  $\Sigma\pi\pi$  (mainly  $\Lambda(1405)\pi$ ) branching fraction produced peripherally, and the other with a large  $\Sigma\pi$  branching fraction produced at larger angles. The experimental results have been confirmed by AGUILAR-BENITEZ 70, ASPELL 74, ESTES 74, and TIMMERMANS 76. If, in fact, there are two resonances, the most likely quantum numbers for both the  $\Sigma\pi$  and the  $\Lambda(1405)\pi$  states are  $D_{13}$ . There is also possibly a third  $\Sigma$  in this region, the  $\Sigma(1690)$  in the Listings, the main evidence for which is a large  $\Lambda\pi/\Sigma\pi$  branching ratio. These topics have been reviewed by EBERHARD 73 and by MILLER 70.

**Formation experiments:** Two states are also observed near this mass in formation experiments. One of these, the  $\Sigma(1670)D_{13}$ , has the same quantum numbers as those observed in production and has a large  $\Sigma\pi/\Sigma\pi\pi$  branching ratio; it may well be the  $\Sigma(1670)$  produced at larger angles (see TIMMERMANS 76). The other state, the  $\Sigma(1660)P_{11}$ , has different quantum numbers, its  $\Sigma\pi/\Sigma\pi\pi$  branching ratio is unknown, and its relation to the produced  $\Sigma(1670)$  states is obscure.

See key on page 601

# Baryon Particle Listings

## $\Sigma(1670)$

$\Sigma(1670) \ 3/2^-$

$$I(J^P) = 1(\frac{3}{2}^-) \text{ Status: } ****$$

For most results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

Results from production experiments are listed separately in the next entry.

### $\Sigma(1670)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1665 to 1685 (<math>\approx 1670</math>) OUR ESTIMATE</b>			
1678 $\pm 2$	ZHANG	13A	DPWA Multichannel
1673 $\pm 1$	GAO	12	DPWA $\overline{K}N \rightarrow \Lambda\pi$
1665.1 $\pm 4.1$	KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$
1682 $\pm 5$	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$
1679 $\pm 10$	ALSTON-...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$
1670 $\pm 5$	GOPAL	77	DPWA $\overline{K}N$ multichannel
1670 $\pm 6$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$
1685 $\pm 20$	BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$
1659 $+12$ $-5$	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
1670 $\pm 2$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1667 or 1668	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
1650	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$
1671 $\pm 3$	PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$ (sol. 1)
1655 $\pm 2$	PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$ (sol. 2)

### $\Sigma(1670)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>40 to 80 (<math>\approx 60</math>) OUR ESTIMATE</b>			
55 $\pm 4$	ZHANG	13A	DPWA Multichannel
52 $+5$ $-2$	GAO	12	DPWA $\overline{K}N \rightarrow \Lambda\pi$
65.0 $\pm 7.3$	KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$
79 $\pm 10$	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$
56 $\pm 20$	ALSTON-...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$
50 $\pm 5$	GOPAL	77	DPWA $\overline{K}N$ multichannel
56 $\pm 3$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$
85 $\pm 25$	BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$
32 $\pm 11$	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
79 $\pm 6$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
46 or 46	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
80	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$
44 $\pm 11$	PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$ (sol. 1)
76 $\pm 5$	PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$ (sol. 2)

### $\Sigma(1670)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1674	ZHANG	13A	DPWA Multichannel

#### $-2\times$ IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
54	ZHANG	13A	DPWA Multichannel

### $\Sigma(1670)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \ N\overline{K}$	7–13 %
$\Gamma_2 \ \Lambda\pi$	5–15 %
$\Gamma_3 \ \Sigma\pi$	30–60 %
$\Gamma_4 \ \Lambda\pi\pi$	
$\Gamma_5 \ \Sigma\pi\pi$	
$\Gamma_6 \ \Sigma(1385)\pi$	
$\Gamma_7 \ \Sigma(1385)\pi, \ S\text{-wave}$	
$\Gamma_8 \ \Lambda(1405)\pi$	
$\Gamma_9 \ \Lambda(1520)\pi$	

The above branching fractions are our estimates, not fits or averages.

### $\Sigma(1670)$ BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
<b>0.07 to 0.13 OUR ESTIMATE</b>				
0.062 $\pm 0.007$	ZHANG	13A	DPWA Multichannel	
0.10 $\pm 0.03$	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$	
0.11 $\pm 0.03$	ALSTON-...	78	DPWA $\overline{K}N \rightarrow \overline{K}N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.08 $\pm 0.03$	GOPAL	77	DPWA See GOPAL 80	
0.07 or 0.07	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1670) \rightarrow \Lambda\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
<b>VALUE</b>				
+0.08 $\pm 0.01$	ZHANG	13A	DPWA Multichannel	
+0.081 $+0.002$ $-0.004$	GAO	12	DPWA $\overline{K}N \rightarrow \Lambda\pi$	
+0.17 $\pm 0.03$	<sup>2</sup> MORRIS	78	DPWA $K^-n \rightarrow \Lambda\pi^-$	
+0.13 $\pm 0.02$	<sup>2</sup> MORRIS	78	DPWA $K^-n \rightarrow \Lambda\pi^-$	
+0.10 $\pm 0.02$	GOPAL	77	DPWA $\overline{K}N$ multichannel	
+0.06 $\pm 0.02$	BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$	
+0.09 $\pm 0.02$	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$	
+0.018 $\pm 0.060$	DEVENISH	74B	Fixed- $t$ dispersion rel.	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
+0.08 or +0.08	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	
+0.05	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$	
+0.08 $\pm 0.01$	PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$ (sol. 1)	
+0.17 $\pm 0.01$	PONTE	75	DPWA $K^-p \rightarrow \Lambda\pi^0$ (sol. 2)	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1670) \rightarrow \Sigma\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
<b>VALUE</b>				
+0.20 $\pm 0.01$	ZHANG	13A	DPWA Multichannel	
+0.20 $\pm 0.02$	KOISO	85	DPWA $K^-p \rightarrow \Sigma\pi$	
+0.21 $\pm 0.02$	GOPAL	77	DPWA $\overline{K}N$ multichannel	
+0.20 $\pm 0.01$	HEPP	76B	DPWA $K^-N \rightarrow \Sigma\pi$	
+0.21 $\pm 0.03$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
+0.18 or +0.17	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	

$\Gamma(\Lambda\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_4/\Gamma$
<b>VALUE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.11	ARMENTEROS68E	HBC	$K^-p$ ( $\Gamma_1=0.09$ )	

$\Gamma(\Sigma\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_5/\Gamma$
<b>VALUE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.14	<sup>3</sup> ARMENTEROS68E	HBC	$K^-p, K^-d$ ( $\Gamma_1=0.09$ )	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1670) \rightarrow \Sigma(1385)\pi, \ S\text{-wave}$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_7)^{1/2}/\Gamma$
<b>VALUE</b>				
+0.11 $\pm 0.03$	PREVOST	74	DPWA $K^-N \rightarrow \Sigma(1385)\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.17 $\pm 0.02$	<sup>4</sup> SIMS	68	DBC $K^-N \rightarrow \Lambda\pi\pi$	

$\Gamma(\Lambda(1405)\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma$
<b>VALUE</b>				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.06	ARMENTEROS68E	HBC	$K^-p, K^-d$ ( $\Gamma_1=0.09$ )	

$\Gamma_1\Gamma_f/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1670) \rightarrow \Lambda(1405)\pi$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1\Gamma_8/\Gamma^2$
<b>VALUE</b>				
0.007 $\pm 0.002$	<sup>5</sup> BRUCKER	70	DBC $K^-N \rightarrow \Sigma\pi\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.03	BERLEY	69	HBC $K^-p$ 0.6–0.82 GeV/ $c$	

$\Gamma(\Lambda(1405)\pi)/\Gamma(\Sigma(1385)\pi)$	DOCUMENT ID	TECN	COMMENT	$\Gamma_8/\Gamma_6$
<b>VALUE</b>				
0.23 $\pm 0.08$	BRUCKER	70	DBC $K^-N \rightarrow \Sigma\pi\pi$	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1670) \rightarrow \Lambda(1520)\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_9)^{1/2}/\Gamma$
<b>VALUE</b>				
0.081 $\pm 0.016$	<sup>6</sup> CAMERON	77	DPWA $P$ -wave decay	

# Baryon Particle Listings

## $\Sigma(1670)$ , $\Sigma(1670)$ Bumps

### $\Sigma(1670)$ FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> Results are with and without an  $S_{11}$   $\Sigma(1620)$  in the fit.  
<sup>3</sup> Ratio only for  $\Sigma 2\pi$  system in  $l = 1$ , which cannot be  $\Sigma(1385)$ .  
<sup>4</sup> SIMS 68 uses only cross-section data. Result used as upper limit only.  
<sup>5</sup> Assuming the  $\Lambda(1405)\pi$  cross-section bump is due only to  $3/2^-$  resonance.  
<sup>6</sup> The CAMERON 77 upper limit on  $F$ -wave decay is 0.03.

### $\Sigma(1670)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GAO	12	PR C86 025201	P. Gao, J. Shi, B.S. Zou	(BHEP, BEIJT)
Also		NP A867 41	P. Gao, B.S. Zou, A. Sibirtsev	(BHEP, BEIJT+)
KOISO	85	NP A433 619	H. Koiso <i>et al.</i>	(TOKY, MASA)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HEL5, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
MORRIS	78	PR D17 55	W.A. Morris <i>et al.</i>	(FSU) IJP
CAMERON	77	NP B131 399	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pldcock, R.G. Moorhouse	(LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pldcock	(LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Pldcock	(LOUC) IJP
DEBELLEFON	76	NP B109 129	A. de Bellefon, A. Berthon	(CDEF) IJP
HEPP	76B	PL 65B 487	V. Hepp <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
PONTE	75	PR D12 2597	R.A. Ponte <i>et al.</i>	(MASA, TENN, UCR) IJP
VANHORN	75	NP B87 145	A.J. van Horn	(LBL) IJP
Also		NP B87 157	A.J. van Horn	(LBL) IJP
DEVENISH	74B	NP B81 330	R.C. Devenish, C.D. Froggatt, B.R. Martin	(DESY+) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i>	(SACL, CERN, HEID)
BRUCKER	70	Duke Conf. 155	E.B. Brucker <i>et al.</i>	(FSU) I
Hyperon Resonances, 1970				
BERLEY	69	PL 30B 430	D. Berley <i>et al.</i>	(BNL)
ARMENTEROS	68E	PL 28B 521	R. Armenteros <i>et al.</i>	(CERN, HEID, SACL) I
SIMS	68	PRL 21 1413	W.H. Sims <i>et al.</i>	(FSU, TUFTS, BRAN)

## $\Sigma(1670)$ Bumps

$$I(J^P) = 1(?)^?$$

### OMITTED FROM SUMMARY TABLE

Formation experiments are listed separately in the preceding entry.

Probably there are two states at the same mass with the same quantum numbers, one decaying to  $\Sigma\pi$  and  $\Lambda\pi$ , the other to  $\Lambda(1405)\pi$ . See the note in front of the preceding entry.

### $\Sigma(1670)$ MASS (PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
≈ 1670 OUR ESTIMATE					
1670 ± 4		<sup>1</sup> CARROLL	76	DPWA	Isospin-1 total σ
1675 ± 10		<sup>2</sup> HEPP	76	DBC	− K <sup>−</sup> N 1.6–1.75 GeV/c
1665 ± 1		APSELL	74	HBC	K <sup>−</sup> p 2.87 GeV/c
1688 ± 2 or 1683 ± 5	1.2k	BERTHON	74	HBC	0 Quasi-2-body σ
1670 ± 6		AGUILAR-...	70B	HBC	K <sup>−</sup> p → Σππ 4 GeV
1668 ± 10		AGUILAR-...	70B	HBC	K <sup>−</sup> p → Σ3π 4 GeV
1660 ± 10		ALVAREZ	63	HBC	+ K <sup>−</sup> p 1.51 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1668 ± 10	150	<sup>3</sup> FERRERSORIA	81	OMEG	− π <sup>−</sup> p 9,12 GeV/c
1655 to 1677		TIMMERMANS	76	HBC	+ K <sup>−</sup> p 4.2 GeV/c
1665 ± 5		BUGG	68	CNTR	K <sup>−</sup> p, d total σ
1661 ± 9	70	PRIMER	68	HBC	+ See BARNES 69E
1685		ALEXANDER	62c	HBC	−0 π <sup>−</sup> p 2–2.2 GeV/c

### $\Sigma(1670)$ WIDTH (PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
67.0 ± 2.4		APSELL	74	HBC	$K^- p$ 2.87 GeV/c	
110 ± 12		AGUILAR-...	70B	HBC	$K^- p \rightarrow \Sigma \pi \pi$ 4 GeV	
135 <sup>+40</sup> <sub>-30</sub>		AGUILAR-...	70B	HBC	$K^- p \rightarrow \Sigma 3 \pi$ 4 GeV	
40 ± 10		ALVAREZ	63	HBC	+	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
90 ± 20	150	<sup>3</sup> FERRERSORIA	81	OMEG	$\pi^- p$ 9,12 GeV/c	
52		<sup>1</sup> CARROLL	76	DPWA	Isospin-1 total $\sigma$	
48 to 63		TIMMERMANS	76	HBC	+	$K^- p$ 4.2 GeV/c
30 ± 15		BUGG	68	CNTR		
60 ± 20	70	PRIMER	68	HBC	+	See BARNES 69E
45		ALEXANDER	62c	HBC	-0	

### $\Sigma(1670)$ DECAY MODES (PRODUCTION EXPERIMENTS)

Mode	
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Lambda\pi$
$\Gamma_3$	$\Sigma\pi$
$\Gamma_4$	$\Lambda\pi\pi$
$\Gamma_5$	$\Sigma\pi\pi$
$\Gamma_6$	$\Sigma(1385)\pi$
$\Gamma_7$	$\Lambda(1405)\pi$

### $\Sigma(1670)$ BRANCHING RATIOS (PRODUCTION EXPERIMENTS)

$\Gamma(N\bar{K})/\Gamma(\Sigma\pi)$					$\Gamma_1/\Gamma_3$	
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
<0.03		TIMMERMANS	76	HBC	+	$K^- p$ 4.2 GeV/c
<0.10		BERTHON	74	HBC	0	Quasi-2-body $\sigma$
<0.2		AGUILAR...	70B	HBC		
<0.26		BARNES	69E	HBC	+	$K^- p$ 3.9-5 GeV/c
0.025		BUGG	68	CNTR	0	Assuming $J = 3/2$
<0.24	0	PRIMER	68	HBC	+	$K^- p$ 4.6-5 GeV/c
<0.6		LONDON	66	HBC	+	$K^- p$ 2.25 GeV/c
<0.19	0	ALVAREZ	63	HBC	+	$K^- p$ 1.15 GeV/c
$\geq 0.5 \pm 0.25$		SMITH	63	HBC	-0	

$\Gamma(\Lambda\pi)/\Gamma(\Sigma\pi)$					$\Gamma_2/\Gamma_3$	
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
0.76 $\pm$ 0.09		ESTES	74	HBC	0	$K^- p$ 2.1,2.6 GeV/c
0.45 $\pm$ 0.15		BARNES	69E	HBC	+	$K^- p$ 3.9-5 GeV/c
0.15 $\pm$ 0.07		HUWE	69	HBC	+	
0.11 $\pm$ 0.06	33	BUTTON-...	68	HBC	+	$K^- p$ 1.7 GeV/c
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
$\leq 0.45 \pm 0.07$		TIMMERMANS	76	HBC	+	$K^- p$ 4.2 GeV/c
0.55 $\pm$ 0.11		BERTHON	74	HBC	0	Quasi-2-body $\sigma$
0	0	PRIMER	68	HBC	+	See BARNES 69E
<0.6		LONDON	66	HBC	+	$K^- p$ 2.25 GeV/c
1.2	130	ALVAREZ	63	HBC	+	$K^- p$ 1.15 GeV/c
1.2		SMITH	63	HBC	-0	

$\Gamma(\Lambda\pi\pi)/\Gamma(\Sigma\pi)$					$\Gamma_4/\Gamma_3$	
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
<0.6	90	LONDON	66	HBC	+	$K^- p$ 2.25 GeV/c
0.56		ALVAREZ	63	HBC	+	$K^- p$ 1.15 GeV/c
0.17		SMITH	63	HBC	−0	

$\Gamma(\Sigma\pi\pi)/\Gamma(\Sigma\pi)$					$\Gamma_5/\Gamma_3$	
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
largest at small angles		ESTES	74	HBC	0	$K^-p$ 2.1,2.6 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.2		<sup>2</sup> HEPP	76	DBC	−	$K^-N$ 1.6–1.75 GeV/c
0.56	180	ALVAREZ	63	HBC	+	$K^-p$ 1.15 GeV/c

$\Gamma(\Lambda(1405)\pi)/\Gamma(\Sigma\pi)$					$\Gamma_7/\Gamma_3$	
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
1.8 $\pm$ 0.3 to 0.02 $\pm$ 0.07		3,4 TIMMERMANS76	HBC	+	$K^- p$ 4.2 GeV/c	
largest at small angles		ESTES	74	HBC	$\pm$ $K^- p$ 2.1, 2.6 GeV/c	
3.0 $\pm$ 1.6	50	LONDON	66	HBC	+	$K^- p$ 2.25 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.58 $\pm$ 0.20	17	PRIMER	68	HBC	+	See BARNES 69E

$\Gamma(\Sigma\pi)/\Gamma(\Sigma\pi\pi)$	$\Gamma_3/\Gamma_5$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
varies with prod. angle	<sup>5</sup> APSELL 74	HBC	+	$K^-p$ 2.87 GeV/c	
1.39±0.16	BERTHON 74	HBC	0	Quasi-2-body $\sigma$	
2.5 to 0.24	<sup>4</sup> EBERHARD 69	HBC		$K^-p$ 2.6 GeV/c	
<0.4	BIRMINGHAM 66	HBC	+	$K^-p$ 3.5 GeV/c	
0.30±0.15	LONDON 66	HBC	+	$K^-p$ 2.25 GeV/c	

$\Gamma(\Lambda(1405)\pi)/\Gamma(\Sigma\pi\pi)$					$\Gamma_7/\Gamma_5$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
$0.97 \pm 0.08$	TIMMERMANS	76	HBC	$K^- p$ 4.2 GeV/c	
$1.00 \pm 0.02$	APSELL	74	HBC	$K^- p$ 2.87 GeV/c	
$0.90^{+0.10}_{-0.16}$	EBERHARD	65	HBC	+	$K^- p$ 2.45 GeV/c

$\Gamma(\Lambda(1405)\pi)/\Gamma(\Sigma(1385)\pi)$	$\Gamma_7/\Gamma_6$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<0.8	EBERHARD	65	HBC	+	$K^- p$ 2.45 GeV/c

See key on page 601

## Baryon Particle Listings

 $\Sigma(1670)$  Bumps,  $\Sigma(1690)$  Bumps,  $\Sigma(1730)$ 

$\Gamma(\Lambda\pi)/\Gamma(\Sigma\pi\pi)$					$\Gamma_4/\Gamma_5$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
$0.35 \pm 0.2$	BIRMINGHAM 66	HBC	+	$K^- p$ 3.5 GeV/c					
$\Gamma(\Lambda\pi)/\Gamma(\Sigma\pi\pi)$					$\Gamma_2/\Gamma_5$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<0.2	BIRMINGHAM 66	HBC	+	$K^- p$ 3.5 GeV/c					
$\Gamma(\Lambda\pi)/[\Gamma(\Lambda\pi) + \Gamma(\Sigma\pi)]$					$\Gamma_2/(\Gamma_2 + \Gamma_3)$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<0.6	AGUILAR-... 70B	HBC							
$\Gamma(\Sigma(1385)\pi)/\Gamma(\Sigma\pi)$					$\Gamma_6/\Gamma_3$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
$\leq 0.21 \pm 0.05$	TIMMERMAN576	HBC		$K^- p$ 4.2 GeV/c					

 $\Sigma(1670)$  QUANTUM NUMBERS  
(PRODUCTION EXPERIMENTS)

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$J^P = 3/2^-$	400	BUTTON-... 68	HBC	$\pm$	$\Sigma^0 \pi$
$J^P = 3/2^-$		EBERHARD 67	HBC	+	$\Lambda(1405) \pi$
$J^P = 3/2^+$		LEVEQUE 65	HBC		$\Lambda(1405) \pi$

 $\Sigma(1670)$  FOOTNOTES

- <sup>1</sup> Total cross-section bump with  $(J+1/2) \Gamma_{\text{el}} / \Gamma_{\text{total}} = 0.23$ .  
<sup>2</sup> Enhancements in  $\Sigma \pi$  and  $\Sigma \pi \pi$  cross sections.  
<sup>3</sup> Backward production in the  $\Lambda \pi^- K^+$  final state.  
<sup>4</sup> Depending on production angle.  
<sup>5</sup> APSELL 74, ESTES 74, and TIMMERMAN576 find strong branching ratio dependence on production angle, as in earlier production experiments.

 $\Sigma(1670)$  REFERENCES  
(PRODUCTION EXPERIMENTS)

FERRERSORIA 81	NP B178 373	A. Ferrer Soria <i>et al.</i>	(CERN, CDEF, EPOL+)
CARROLL 76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL) I
HEPP 76	NP B115 82	V. Hepp <i>et al.</i>	(CERN, HEID, MPIM) I
TIMMERMAN576	NP B112 77	J.J.M. Timmermans <i>et al.</i>	(NUM, CERN+) JP
APSELL 74	PR D10 1419	S.P. Appell <i>et al.</i>	(BRAN, UMD, SYRA+) I
BERTHON 74	NC 21A 146	A. Berthon <i>et al.</i>	(CDEF, RHEL, SACL+) I
ESTES 74	Thesis LBL-3827	R.D. Estes	(LBL)
AGUILAR-... 70B	PRL 25 58	M. Aguilar-Benitez <i>et al.</i>	(BNL, SYRA)
BARNES 69E	BNL 13823	V.E. Barnes <i>et al.</i>	(BNL, SYRA)
EBERHARD 69	PRL 22 200	P.H. Eberhard <i>et al.</i>	(LRL)
HUVE 69	PR 181 1824	D.O. Huve	(LRL)
BUGG 68	PR 168 1466	D.V. Bugg <i>et al.</i>	(RHEL, BIRM, CAVE) I
BUTTON-... 68	PRL 21 1123	J. Button-Shafer	(MASA, LRL) JP
PRIMER 68	PRL 20 610	M. Primer <i>et al.</i>	(SYRA, BNL) I
EBERHARD 67	PR 163 1446	P. Eberhard <i>et al.</i>	(LRL, ILL) JP
BIRMINGHAM 66	PR 152 1148	M. Hogue <i>et al.</i>	(BIRM, GLAS, LOIC, OXF+) I
LONDON 66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA) IJ
EBERHARD 65	PRL 14 466	P.H. Eberhard <i>et al.</i>	(LRL, ILL) I
LEVEQUE 65	PL 18 69	A. Leveque <i>et al.</i>	(SACL, EPOL, GLAS+) JP
ALVAREZ 63	PRL 10 184	L.W. Alvarez <i>et al.</i>	(LRL) I
SMITH 63	Athens Conf. 67	G.A. Smith	(LRL)
ALEXANDER 62C	CERN Conf. 320	G. Alexander <i>et al.</i>	(LRL) I

 $\Sigma(1690)$  Bumps

$$I(J^P) = 1(?)^? \quad \text{Status: **}$$

OMITTED FROM SUMMARY TABLE

See the note preceding the  $\Sigma(1670)$  Listings. Seen in production experiments only, mainly in  $\Lambda \pi$ . $\Sigma(1690)$  MASS  
(PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>\approx 1690</math> OUR ESTIMATE</b>					
$1698 \pm 20$	70	<sup>1</sup> GODDARD 79	HBC	+	$\pi^+ p$ 10.3 GeV/c
$1707 \pm 20$	40	<sup>2</sup> GODDARD 79	HBC	+	$\pi^+ p$ 10.3 GeV/c
$1698 \pm 20$	15	ADERHOLZ 69	HBC	+	$\pi^+ p$ 8 GeV/c
$1682 \pm 2$	46	BLUMENFELD 69	HBC	+	$K_L^0 p$
$1700 \pm 20$		MOTT 69	HBC	+	$K^- p$ 5.5 GeV/c
$1694 \pm 24$	60	<sup>3</sup> PRIMER 68	HBC	+	$K^- p$ 4.6-5 GeV/c
$1700 \pm 6$		<sup>4</sup> SIMS 68	HBC	-	$K^- N \rightarrow \Lambda \pi \pi$
$1715 \pm 12$	30	COLLEY 67	HBC	+	$K^- p$ 6 GeV/c

 $\Sigma(1690)$  WIDTH  
(PRODUCTION EXPERIMENTS)

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$240 \pm 60$	70	<sup>1</sup> GODDARD 79	HBC	+	$\pi^+ p$ 10.3 GeV/c
$130^{+100}_{-60}$	40	<sup>2</sup> GODDARD 79	HBC	+	$\pi^+ p$ 10.3 GeV/c
$142 \pm 40$	15	ADERHOLZ 69	HBC	+	$\pi^+ p$ 8 GeV/c
$25 \pm 10$	46	BLUMENFELD 69	HBC	+	$K_L^0 p$

$130 \pm 25$		MOTT 69	HBC	+	$K^- p$ 5.5 GeV/c
$105 \pm 35$	60	<sup>3</sup> PRIMER 68	HBC	+	$K^- p$ 4.6-5 GeV/c
$62 \pm 14$		<sup>4</sup> SIMS 68	HBC	-	$K^- N \rightarrow \Lambda \pi \pi$
$100 \pm 35$	30	COLLEY 67	HBC	+	$K^- p$ 6 GeV/c

 $\Sigma(1690)$  DECAY MODES  
(PRODUCTION EXPERIMENTS)

Mode	DOCUMENT ID	TECN	CHG	COMMENT
$\Gamma_1$ $N \bar{K}$				
$\Gamma_2$ $\Lambda \pi$				
$\Gamma_3$ $\Sigma \pi$				
$\Gamma_4$ $\Sigma(1385) \pi$				
$\Gamma_5$ $\Lambda \pi \pi$ (including $\Sigma(1385) \pi$ )				

 $\Sigma(1690)$  BRANCHING RATIOS  
(PRODUCTION EXPERIMENTS)

$\Gamma(N\bar{K})/\Gamma(\Lambda\pi)$					$\Gamma_1/\Gamma_2$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
small		GODDARD 79	HBC	+	$\pi^+p$ 10.2 GeV/c
<0.2		MOTT 69	HBC	+	$K^-p$ 5.5 GeV/c
$0.4 \pm 0.25$	18	COLLEY 67	HBC	+	6/30 events

$\Gamma(\Sigma\pi)/\Gamma(\Lambda\pi)$					$\Gamma_3/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
small		GODDARD 79	HBC	+	$\pi^+ p$ 10.2 GeV/c	
<0.4	90	MOTT 69	HBC	+	$K^- p$ 5.5 GeV/c	
$0.3 \pm 0.3$		COLLEY 67	HBC	+	4/30 events	

$\Gamma(\Sigma(1385)\pi)/\Gamma(\Lambda\pi)$					$\Gamma_4/\Gamma_2$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<0.5	MOTT 69	HBC	+	$K^- p$ 5.5 GeV/c					

$\Gamma(\Lambda\pi \text{ (including } \Sigma(1385)\pi))/\Gamma(\Lambda\pi)$					$\Gamma_5/\Gamma_2$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
$2.0 \pm 0.6$	BLUMENFELD 69	HBC	+	31/15 events					
$0.5 \pm 0.25$	COLLEY 67	HBC	+	15/30 events					

$\Gamma(\Sigma(1385)\pi)/\Gamma(\Lambda\pi \text{ (including } \Sigma(1385)\pi))$					$\Gamma_4/\Gamma_5$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	VALUE	DOCUMENT ID	TECN	CHG	COMMENT
large	SIMS 68	HBC	-	$K^- N \rightarrow \Lambda \pi \pi$					
small	COLLEY 67	HBC	+	$K^- p$ 6 GeV/c					

 $\Sigma(1690)$  FOOTNOTES  
(PRODUCTION EXPERIMENTS)

- <sup>1</sup> From  $\pi^+ p \rightarrow (\Lambda \pi^+) K^+$ .  $J > 1/2$  is not required by the data.  
<sup>2</sup> From  $\pi^+ p \rightarrow (\Lambda \pi^+) (K \pi)^+$ .  $J > 1/2$  is indicated, but large background precludes a definite conclusion.  
<sup>3</sup> See the  $\Sigma(1670)$  Listings. AGUILAR-BENITEZ 70B with three times the data of PRIMER 68 find no evidence for the  $\Sigma(1690)$ .  
<sup>4</sup> This analysis, which is difficult and requires several assumptions and shows no unambiguous  $\Sigma(1690)$  signal, suggests  $J^P = 5/2^+$ . Such a state would lead all previously known  $Y^*$  trajectories.

 $\Sigma(1690)$  REFERENCES  
(PRODUCTION EXPERIMENTS)

GODDARD 79	PR D19 1350	M.C. Goddard <i>et al.</i>	(TNT0, BNL) IJ
AGUILAR-... 70B	PRL 25 58	M. Aguilar-Benitez <i>et al.</i>	(BNL, SYRA)
ADERHOLZ 69	NP B11 259	M. Aderholz <i>et al.</i>	(AACH3, BERL, CERN+) I
BLUMENFELD 69	PL 29B 58	B.J. Blumenfeld, G.R. Kalbfleisch	(BNL) I
MOTT 69	PR 177 1966	J. Mott <i>et al.</i>	(NWES, ANL) I
Also	PRL 18 266	M. Derrick <i>et al.</i>	(ANL, NWES) I
PRIMER 68	PRL 20 610	M. Primer <i>et al.</i>	(SYRA, BNL) I
SIMS 68	PRL 21 1413	W.H. Sims <i>et al.</i>	(FSU, TUFTS, BRAN) I
COLLEY 67	PL 24B 489	D.C. Colley	(BIRM, GLAS, LOIC, MUNI, OXF+) I

 $\Sigma(1730)$   $3/2^+$ 

$$I(J^P) = 1(\frac{3}{2}^+) \quad \text{Status: *}$$

OMITTED FROM SUMMARY TABLE

 $\Sigma(1730)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>1727 \pm 27</math></b>	ZHANG 13A	DPWA	Multichannel

 $\Lambda(1730)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>276 \pm 87</math></b>	ZHANG 13A	DPWA	Multichannel

# Baryon Particle Listings

## $\Sigma(1730)$ , $\Sigma(1750)$

Σ(1730) DECAY MODES				
Mode	Fraction (Γ <sub>i</sub> /Γ)			
Γ <sub>1</sub>	N $\overline{K}$	( 2.0 ± 1.0 ) %		
Γ <sub>2</sub>	Λπ	( 70 ±17 ) %		
Γ <sub>3</sub>	Σπ	( 12 ± 6 ) %		

Σ(1730) BRANCHING RATIOS				
Γ(N $\overline{K}$ )/Γ <sub>total</sub>	Γ <sub>1</sub> /Γ			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.02±0.01	ZHA NG	13A	DPWA	Multichannel
Γ(Λπ)/Γ <sub>total</sub>	Γ <sub>2</sub> /Γ			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.70±0.17	ZHA NG	13A	DPWA	Multichannel
Γ(Σπ)/Γ <sub>total</sub>	Γ <sub>3</sub> /Γ			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.12±0.06	ZHA NG	13A	DPWA	Multichannel

Σ(1730) REFERENCES				
ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)

Σ(1750) 1/2<sup>−</sup>

I(J<sup>P</sup>) = 1(½<sup>−</sup>) Status: \*\*\*

For most results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

There is evidence for this state in many partial-wave analyses, but with wide variations in the mass, width, and couplings. The latest analyses indicated significant couplings to N $\overline{K}$  and Λπ, as well as to Ση whose threshold is at 1746 MeV (JONES 74).

Σ(1750) MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>1730 to 1800 (≈ 1750) OUR ESTIMATE</b>				
1739± 8	ZHA NG	13A	DPWA	Multichannel
1756±10	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1770±10	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1770±15	GOPAL	77	DPWA	$\overline{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1800 or 1813	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
1715±10	<sup>2</sup> CARROLL	76	DPWA	Isospin-1 total σ
1730	DEBELLEFON	76	IPWA	$K^-p \rightarrow \Lambda\pi^0$
1780±30	BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$ (sol. 1)
1700±30	BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$ (sol. 2)
1697 <sup>+20</sup> <sub>−10</sub>	VANHORN	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
1785±12	CHU	74	DBC	Fits σ(K <sup>−</sup> n → Σ <sup>−</sup> η)
1760± 5	<sup>3</sup> JONES	74	HBC	Fits σ(K <sup>−</sup> p → Σ <sup>0</sup> η)
1739±10	PREVOST	74	DPWA	K <sup>−</sup> N → Σ(1385) π

Σ(1750) WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
<b>60 to 160 (≈ 90) OUR ESTIMATE</b>				
182±60	ZHA NG	13A	DPWA	Multichannel
64±10	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
161±20	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
60±10	GOPAL	77	DPWA	$\overline{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
117 or 119	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
10	<sup>2</sup> CARROLL	76	DPWA	Isospin-1 total σ
110	DEBELLEFON	76	IPWA	K <sup>−</sup> p → Λπ <sup>0</sup>
140±30	BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$ (sol. 1)
160±50	BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$ (sol. 2)
66 <sup>+14</sup> <sub>−12</sub>	VANHORN	75	DPWA	K <sup>−</sup> p → Λπ <sup>0</sup>
89±33	CHU	74	DBC	Fits σ(K <sup>−</sup> n → Σ <sup>−</sup> η)
92± 7	<sup>3</sup> JONES	74	HBC	Fits σ(K <sup>−</sup> p → Σ <sup>0</sup> η)
108±20	PREVOST	74	DPWA	K <sup>−</sup> N → Σ(1385) π

Σ(1750) POLE POSITION				
REAL PART				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1708	ZHA NG	13A	DPWA	Multichannel

### −2×IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
158	ZHANG	13A	DPWA Multichannel

### $\Sigma(1750)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\overline{K}$	10–40 %
$\Gamma_2$ $\Lambda\pi$	seen
$\Gamma_3$ $\Sigma\pi$	<8 %
$\Gamma_4$ $\Sigma\eta$	15–55 %
$\Gamma_5$ $\Sigma(1385)\pi$ , $D$ -wave	
$\Gamma_6$ $\Lambda(1520)\pi$	
The above branching fractions are our estimates, not fits or averages.	
$\Gamma_7$ $N\overline{K}^*(892)$ , $S=1/2$	(8±4) %

### $\Sigma(1750)$ BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.1 to 0.4 OUR ESTIMATE</b>				
0.09 ± 0.07	ZHANG	13A	DPWA	Multichannel
0.14 ± 0.03	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
0.33 ± 0.05	ALSTON-...	78	DPWA	$\overline{K}N \rightarrow \overline{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.15 ± 0.03	GOPAL	77	DPWA	See GOPAL 80
0.06 or 0.05	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1750) \rightarrow \Lambda\pi$	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.10 ±0.04	ZHANG	13A	DPWA	Multichannel
0.04 ±0.03	GOPAL	77	DPWA	$\overline{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.10 or −0.09	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
−0.12	DEBELLEFON	76	IPWA	$K^-p \rightarrow \Lambda\pi^0$
−0.12 ±0.02	BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$ (sol. 1)
−0.13 ±0.03	BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$ (sol. 2)
−0.13 ±0.04	VANHORN	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
−0.120±0.077	DEVENISH	74B	Fixed- $t$ dispersion rel.	

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1750) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.17±0.07	ZHANG	13A	DPWA	Multichannel
−0.09±0.05	GOPAL	77	DPWA	$\overline{K}N$ multichannel
• • • We do not use the following data for averages, fits, limits, etc. • • •				
+0.06 or +0.06	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
0.13±0.02	LANGBEIN	72	IPWA	$\overline{K}N$ multichannel

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1750) \rightarrow \Sigma\eta$	$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.23±0.01	<sup>3</sup> JONES	74	HBC	Fits $\sigma(K^-p \rightarrow \Sigma^0\eta)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	CLINE	69	DBC	Threshold bump

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1750) \rightarrow \Sigma(1385)\pi$ , $D$ -wave	$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.17±0.07	ZHANG	13A	DPWA	Multichannel
+0.18±0.15	PREVOST	74	DPWA	$K^-N \rightarrow \Sigma(1385)\pi$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1750) \rightarrow \Lambda(1520)\pi$	$(\Gamma_1\Gamma_6)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.032 \pm 0.021$	CAMERON	77	DPWA	$P$ -wave decay

$\Gamma(N\overline{K}^*(892), S=1/2)/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.08±0.04</b>	ZHANG	13A	DPWA	Multichannel

### $\Sigma(1750)$ FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> A total cross-section bump with  $(J+1/2)\Gamma_{\text{el}}/\Gamma_{\text{total}} = 0.30$ .  
<sup>3</sup> An  $S$ -wave Breit-Wigner fit to the threshold cross section with no background and errors statistical only.



See key on page 601

# Baryon Particle Listings

## $\Sigma(1750)$ , $\Sigma(1770)$ , $\Sigma(1775)$

### $\Sigma(1750)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HELSE, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-...	78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
CAMERON	77	NP B131 399	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN	77	NP B127 349	B.R. Martin, M.K. Piddcock, R.G. Moorhouse	(LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Piddcock	(LOUC) IJP
Also		NP B126 285	B.R. Martin, M.K. Piddcock	(LOUC) IJP
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL) IJP
DEBELLEFON	76	NP B109 129	A. de Bellefon, A. Berthon	(CDEF) IJP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
VANHORN	75	NP B87 145	A.J. van Horn	(LBL) IJP
Also		NP B87 157	A.J. van Horn	(LBL) IJP
CHU	74	NC 20A 35	R.Y.L. Chu <i>et al.</i>	(PLAT, TUFTS, BRAN) IJP
DEVENISH	74B	NP B81 330	R.C.E. Devenish, C.D. Froggatt, B.R. Martin	(DESY+) IJP
JONES	74	NP B73 141	M.D. Jones	(CHIC) IJP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i>	(SACL, CERN, HEID) IJP
LANGBEIN	72	NP B47 477	W. Langbein, F. Wagner	(MPIM) IJP
CLINE	69	LNC 2 407	D. Cline, R. Laumann, J. Mapp	(WISC) IJP

### $\Sigma(1770)$ $1/2^+$

$$I(J^P) = 1(\frac{1}{2}^+) \text{ Status: } *$$

OMITTED FROM SUMMARY TABLE

Evidence for this state now rests solely on solution 1 of BAILLON 75, (see the footnotes) but the  $\Lambda\pi$  partial-wave amplitudes of this solution are in disagreement with amplitudes from most other  $\Lambda\pi$  analyses. ZHANG 13A finds no evidence for this state.

### $\Sigma(1770)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 1770</math> OUR ESTIMATE</b>			
1738 $\pm$ 10	<sup>1</sup> GOPAL	77	DPWA $\bar{K}N$ multichannel
1770 $\pm$ 20	<sup>2</sup> BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$
1772	<sup>3</sup> KANE	72	DPWA $K^-p \rightarrow \Sigma\pi$

### $\Sigma(1770)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
72 $\pm$ 10	<sup>1</sup> GOPAL	77	DPWA $\bar{K}N$ multichannel
80 $\pm$ 30	<sup>2</sup> BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$
80	<sup>3</sup> KANE	72	DPWA $K^-p \rightarrow \Sigma\pi$

### $\Sigma(1770)$ DECAY MODES

Mode	
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Lambda\pi$
$\Gamma_3$	$\Sigma\pi$

### $\Sigma(1770)$ BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
0.14 $\pm$ 0.04	<sup>1</sup> GOPAL	77	DPWA $\bar{K}N$ multichannel	
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1770) \rightarrow \Lambda\pi$				$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
< 0.04	GOPAL	77	DPWA $\bar{K}N$ multichannel	
-0.08 $\pm$ 0.02	<sup>2</sup> BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$	
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1770) \rightarrow \Sigma\pi$				$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
< 0.04	GOPAL	77	DPWA $\bar{K}N$ multichannel	
-0.108	<sup>3</sup> KANE	72	DPWA $K^-p \rightarrow \Sigma\pi$	

### $\Sigma(1770)$ FOOTNOTES

- <sup>1</sup> Required to fit the isospin-1 total cross section of CARROLL 76 in the  $\bar{K}N$  channel. The addition of new  $K^-p$  polarization and  $K^-n$  differential cross-section data in GOPAL 80 find it to be more consistent with the  $\Sigma(1660) P_{11}$ .  
<sup>2</sup> From solution 1 of BAILLON 75; not present in solution 2.  
<sup>3</sup> Not required in KANE 74, which supersedes KANE 72.

### $\Sigma(1770)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL)
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
CARROLL	76	PRL 37 806	A.S. Carroll <i>et al.</i>	(BNL) IJP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP
KANE	72	PR D5 1583	D.F.J. Kane	(LBL)

### $\Sigma(1775)$ $5/2^-$

$$I(J^P) = 1(\frac{5}{2}^-) \text{ Status: } ****$$

Discovered by GALTIERI 63, this resonance plays the same role as cornerstone for isospin-1 analyses in this region as the  $\Lambda(1820) F_{05}$  does in the isospin-0 channel.

For most results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

### $\Sigma(1775)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1770 to 1780 (<math>\approx 1775</math>) OUR ESTIMATE</b>			
1778 $\pm$ 1	ZHANG	13A	DPWA Multichannel
1778 $\pm$ 5	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
1777 $\pm$ 5	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
1774 $\pm$ 5	GOPAL	77	DPWA $\bar{K}N$ multichannel
1775 $\pm$ 10	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$
1774 $\pm$ 10	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
1772 $\pm$ 6	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1772 or 1777	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel
1765	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$

### $\Sigma(1775)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>105 to 135 (<math>\approx 120</math>) OUR ESTIMATE</b>			
131 $\pm$ 3	ZHANG	13A	DPWA Multichannel
137 $\pm$ 10	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
116 $\pm$ 10	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
130 $\pm$ 10	GOPAL	77	DPWA $\bar{K}N$ multichannel
125 $\pm$ 15	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$
146 $\pm$ 18	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
154 $\pm$ 10	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
102 or 103	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel
120	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$

### $\Sigma(1775)$ POLE POSITION

#### REAL PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1759	ZHANG	13A	DPWA Multichannel

#### -2 $\times$ IMAGINARY PART

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
118	ZHANG	13A	DPWA Multichannel

### $\Sigma(1775)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Lambda\pi$
$\Gamma_3$	$\Sigma\pi$
$\Gamma_4$	$\Sigma(1385)\pi$
$\Gamma_5$	$\Sigma(1385)\pi$ , D-wave
$\Gamma_6$	$\Lambda(1520)\pi$ , P-wave
$\Gamma_7$	$\Sigma\pi\pi$
$\Gamma_8$	$\Delta(1232)\bar{K}$ , D-wave
$\Gamma_9$	$N\bar{K}^*(892)$ , S=1/2
$\Gamma_{10}$	$N\bar{K}^*(892)$ , S=3/2, D-wave

The above branching fractions are our estimates, not fits or averages.

### CONSTRAINED FIT INFORMATION

An overall fit to 7 branching ratios uses 18 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 363.4$  for 14 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	-44			
$x_3$	-23	10		
$x_4$	-23	-32	-4	
$x_6$	-3	1	1	-84
	$x_1$	$x_2$	$x_3$	$x_4$

## Baryon Particle Listings

 $\Sigma(1775)$  $\Sigma(1775)$  BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances. Also, the errors quoted do not include uncertainties due to the parametrization used in the partial-wave analyses and are thus too small.

 $\Gamma(N\bar{K})/\Gamma_{\text{total}}$  $\Gamma_1/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.37 to 0.43 OUR ESTIMATE</b>			
<b>0.421 ± 0.020 OUR FIT</b>	Error includes scale factor of 2.5.		
<b>0.398 ± 0.009 OUR AVERAGE</b>			
0.40 ± 0.01	ZHANG	13A	DPWA Multichannel
0.40 ± 0.02	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
0.37 ± 0.03	ALSTON-...	78	DPWA $\bar{K}N \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.41 ± 0.03	GOPAL	77	DPWA See GOPAL 80
0.37 or 0.36	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

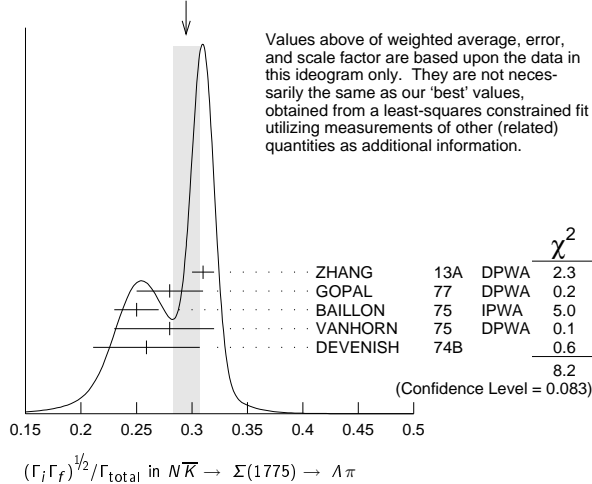
 $\Gamma(\Lambda\pi)/\Gamma(N\bar{K})$  $\Gamma_2/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.48 ± 0.06 OUR FIT</b>	Error includes scale factor of 2.3.		
<b>0.33 ± 0.05</b>	UHLIG	67	HBC $K^- p$ 0.9 GeV/c

 $(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Sigma(1775) \rightarrow \Lambda\pi$  $(\Gamma_1\Gamma_2)^{1/2}/\Gamma$ 

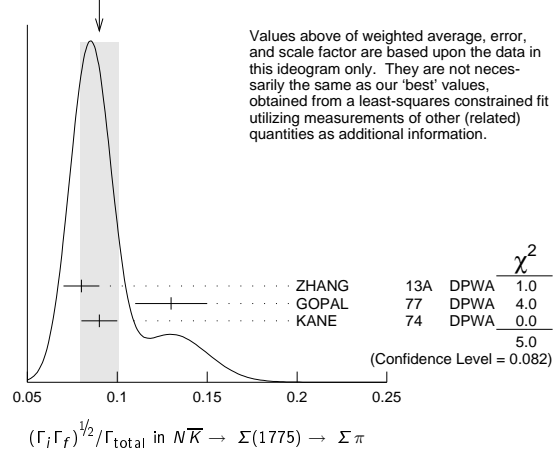
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.293 ± 0.013 OUR FIT</b>	Error includes scale factor of 1.8.		
<b>0.295 ± 0.012 OUR AVERAGE</b>	Signs on measurements were ignored. Error includes scale factor of 1.4. See the ideogram below.		
-0.31 ± 0.01	ZHANG	13A	DPWA Multichannel
-0.28 ± 0.03	GOPAL	77	DPWA $\bar{K}N$ multichannel
-0.25 ± 0.02	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$
-0.28 ± 0.04	VANHORN	75	DPWA $K^- p \rightarrow \Lambda\pi^0$
-0.259 ± 0.048	DEVENISH	74B	Fixed- $t$ dispersion rel.
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.29 or -0.28	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel
-0.30	DEBELLEFON	76	IPWA $K^- p \rightarrow \Lambda\pi^0$

WEIGHTED AVERAGE  
0.295 ± 0.012 (Error scaled by 1.4)

 $(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Sigma(1775) \rightarrow \Sigma\pi$  $(\Gamma_1\Gamma_3)^{1/2}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.090 ± 0.009 OUR FIT</b>	Error includes scale factor of 1.4.		
<b>0.090 ± 0.011 OUR AVERAGE</b>	Signs on measurements were ignored. Error includes scale factor of 1.6. See the ideogram below.		
+0.08 ± 0.01	ZHANG	13A	DPWA Multichannel
+0.13 ± 0.02	GOPAL	77	DPWA $\bar{K}N$ multichannel
0.09 ± 0.01	KANE	74	DPWA $K^- p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+0.08 or +0.08	<sup>1</sup> MARTIN	77	DPWA $\bar{K}N$ multichannel

WEIGHTED AVERAGE  
0.090 ± 0.011 (Error scaled by 1.6)

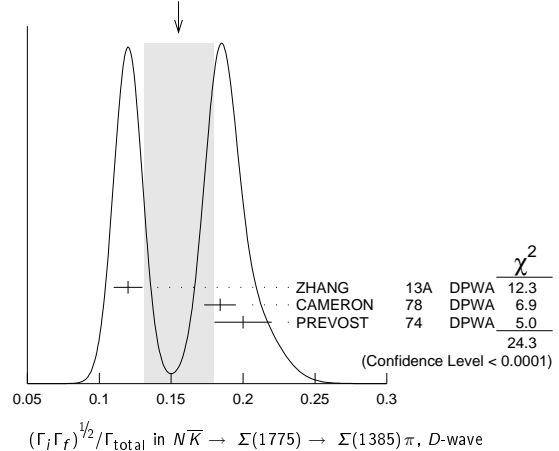
 $\Gamma(\Sigma(1385)\pi)/\Gamma(N\bar{K})$  $\Gamma_4/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.79 ± 0.11 OUR FIT</b>	Error includes scale factor of 3.2.		
<b>0.25 ± 0.09</b>	UHLIG	67	HBC $K^- p$ 0.9 GeV/c

 $(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Sigma(1775) \rightarrow \Sigma(1385)\pi$ , D-wave $(\Gamma_1\Gamma_5)^{1/2}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.155 ± 0.024 OUR AVERAGE</b>	Signs on measurements were ignored. Error includes scale factor of 3.5. See the ideogram below.		
-0.12 ± 0.01	ZHANG	13A	DPWA Multichannel
-0.184 ± 0.011	<sup>2</sup> CAMERON	78	DPWA $K^- p \rightarrow \Sigma(1385)\pi$
+0.20 ± 0.02	PREVOST	74	DPWA $K^- N \rightarrow \Sigma(1385)\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.32 ± 0.06	SIMS	68	DBC $K^- N \rightarrow \Lambda\pi\pi$
0.24 ± 0.03	ARMENTEROS67C	HBC	$K^- p \rightarrow \Lambda\pi\pi$

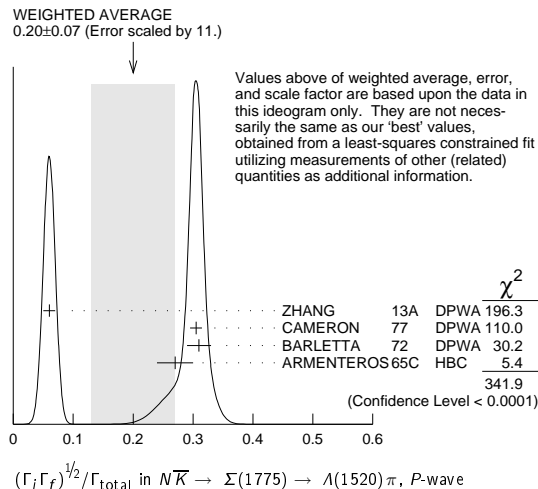
WEIGHTED AVERAGE  
0.155 ± 0.024 (Error scaled by 3.5)

 $\Gamma(\Lambda(1520)\pi, P\text{-wave})/\Gamma(N\bar{K})$  $\Gamma_6/\Gamma_1$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.053 ± 0.080 OUR FIT</b>	Error includes scale factor of 11.8.		
<b>-0.035</b>			
<b>0.28 ± 0.05</b>	UHLIG	67	HBC $K^- p$ 0.9 GeV/c

 $(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Sigma(1775) \rightarrow \Lambda(1520)\pi$ , P-wave $(\Gamma_1\Gamma_6)^{1/2}/\Gamma$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.10 ± 0.06 OUR FIT</b>	Error includes scale factor of 11.5.		
<b>0.20 ± 0.07 OUR AVERAGE</b>	Signs on measurements were ignored. Error includes scale factor of 10.7. See the ideogram below.		
-0.06 ± 0.01	ZHANG	13A	DPWA Multichannel
-0.305 ± 0.010	<sup>3</sup> CAMERON	77	DPWA $K^- p \rightarrow \Lambda(1520)\pi^0$
0.31 ± 0.02	BARLETTA	72	DPWA $K^- p \rightarrow \Lambda(1520)\pi^0$
0.27 ± 0.03	ARMENTEROS65C	HBC	$K^- p \rightarrow \Lambda(1520)\pi^0$



$\Gamma(\Sigma\pi\pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_7/\Gamma$
0.12	4	ARMENTEROS68C	HDBC $K^- N \rightarrow \Sigma\pi\pi$	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1775) \rightarrow \Delta(1232)\bar{K}$ , $D$ -wave	$(\Gamma_1\Gamma_8)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$0.06 \pm 0.03$	ZHA NG	13A	DPWA Multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1775) \rightarrow N\bar{K}^*(892), S=1/2$				$(\Gamma_1\Gamma_g)^{1/2}/\Gamma_{\text{total}}$
VALUE	DUPLICATION ID	TECN	COMMENT	
$+0.04 \pm 0.01$	ZHA NG	13A	DPWA	Multichannel

$(\Gamma_1 \Gamma_r)^{1/2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1775) \rightarrow N\bar{K}^*(892), S=3/2, D\text{-wave}$				
VALUE	DOCUMENT ID	TECN	COMMENT	
$+0.04 \pm 0.01$	ZHANG	13a	DPWA	Multichannel

## Σ(1775) FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> The CAMERON 78 upper limit on G-wave decay is 0.03.  
<sup>3</sup> This rate combines P-wave- and F-wave decays. The CAMERON 77 results for the separate P-wave- and F-wave decays are  $-0.303 \pm 0.010$  and  $-0.037 \pm 0.014$ . The published signs have been changed here to be in accord with the baryon-first convention.  
<sup>4</sup> For about 3/4 of this, the  $\Sigma\pi$  system has  $l=0$  and is almost entirely  $\Lambda(1520)$ . For the rest, the  $\Sigma\pi$  has  $l=1$ , which is about what is expected from the known  $\Sigma(1775) \rightarrow \Sigma(1385)\pi$  rate, as seen in  $\Lambda\pi\pi$ .

## Σ(1775) REFERENCES

ZHANG	132	PR 118 035205	H. Zhang <i>et al.</i>	(KSU)
PDG	82	PL 11B 1	M. Roos <i>et al.</i>	(HELS, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL) JUP
ALSTON....	78	NP D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) JUP
Also		PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) JUP
CAMERON	78	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) JUP
CAMERON	77	NP B131 399	W. Cameron <i>et al.</i>	(RHEL, LOIC) JUP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) JUP
MARTIN	77	NP B127 344	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) JUP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC)
Also		NP B126 265	B.R. Martin, M.K. Pidcock	(LOUC)
DEBELLEFON	76	NP B109 129	A. de Bellefon, A. Berthon	(CDEF) JUP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) JUP
VANHORN	75	NP B87 145	A.J. van Horn	(LBL) JUP
Also		NP B87 157	A.J. van Horn	(LBL) JUP
DEVENISH	74B	NP B81 330	R.C.E. Devenish, C.D. Froggatt, B.R. Martin	(DESY+) JUP
KANE	74	LBL-2452	D.F. Kane	(LBL) JUP
PREVOST	74	NP B69 246	J. Prevost <i>et al.</i>	(SACL, CERN, HEID)
BARLETTA	72	NP B40 45	W.A. Barletta	(EFI) JUP
Also		PRL 17 841	S. Fenster <i>et al.</i>	(CHIC, ANL, CERN) JUP
ARMENTEROS	68C	NP B6 216	R. Armenteros <i>et al.</i>	(CERN, HEID, SL) JUP
SIMS	68	PR 21 1413	W. Sims <i>et al.</i>	(FSU, TUFTS, BRAN)
ARMENTEROS	67C	ZPHY 202 486	R. Armenteros <i>et al.</i>	(CERN, HEID, SACL)
UHLIG	67	NP 155 1448	R.P. Uhlig <i>et al.</i>	(UMD, NRL)
ARMENTEROS	65C	PL 19 338	R. Armenteros <i>et al.</i>	(CERN, HEID, SACL) JUP
GALTIERI	63	PL 6 296	A. Galtieri, A. Hussain, R. Tripp	(LRL) JUP

$\Sigma(1840) \ 3/2^+$

$$I(J^P) = 1(\frac{3}{2}^+) \text{ Status: } *$$

OMITTED FROM SUMMARY TABLE

For the time being, we list together here all resonance claims in the  $P_{13}$  wave between 1700 and 1900 MeV.

**$\Sigma(1840)$  MASS**

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 1840</math> OUR ESTIMATE</b>			
1798 or 1802	<sup>1</sup> MARTIN	77	DPWA $\overline{K} N$ multichannel
1720 $\pm 30$	<sup>2</sup> BAILLON	75	IPWA $\overline{K} N \rightarrow A \pi$
1925 $\pm 200$	VAN HORN	75	DPWA $K^- p \rightarrow A \pi^0$
1840 $\pm 10$	LANGBEIN	72	IPWA $\overline{K} N$ multichannel

**$\Sigma(1840)$  WIDTH**

VALUE (MeV)		DOCUMENT ID	TECN	COMMENT
93 or 93	<sup>1</sup>	MARTIN	77	DPWA $\bar{K} N$ multichannel
120 $\pm$ 30	<sup>2</sup>	BAILLON	75	IPWA $\bar{K} N \rightarrow \Lambda \pi$
65 $^{+50}_{-20}$		VANHORN	75	DPWA $K^- p \rightarrow \Lambda \pi^0$
120 $\pm$ 10		LANGBEIN	72	IPWA $\bar{K} N$ multichannel

## Σ(1840) DECAY MODES

	Mode
$\Gamma_1$	$N\overline{K}$
$\Gamma_2$	$\Lambda\pi$
$\Gamma_3$	$\Sigma\pi$

### $\Sigma(1840)$ BRANCHING RATIOS

See “Sign conventions for resonance couplings” in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(NK)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
0 or 0	<sup>1</sup> MARTIN	77 DPWA	$\overline{KN}$ multichannel	
$0.37 \pm 0.13$	LANGBEIN	72 IPWA	$\overline{KN}$ multichannel	

$(\Gamma_1 \Gamma_2)^{1/2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1840) \rightarrow \Lambda\pi$				$(\Gamma_1 \Gamma_2)^{1/2} / \Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.03 or +0.03	1 MARTIN	77 DPWA	$\bar{K}N$ multichannel	
+0.11 $\pm$ 0.02	2 BAILLON	75 IPWA	$\bar{K}N \rightarrow \Lambda\pi$	
+0.06 $\pm$ 0.04	VANHORN	75 DPWA	$K^- p \rightarrow \Lambda\pi^0$	
+0.122 $\pm$ 0.078	DEVENISH	74B	Fixed- $t$ dispersion rel.	
0.20 $\pm$ 0.04	LANGBEIN	72 IPWA	$\bar{K}N$ multichannel	

$(\Gamma_1 \Gamma_2)^{1/2} / \Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1840) \rightarrow \Sigma \pi$				$(\Gamma_1 \Gamma_3)^{1/2} / \Gamma_{\text{total}}$
VALUE		DOCUMENT ID	TECN	COMMENT
$-0.04$ or $-0.04$	<sup>1</sup>	MARTIN	77	DPWA $\bar{K}N$ multichannel
$0.15 \pm 0.04$		LANGBEIN	72	IPWA $\bar{K}N$ multichannel

## Σ(1840) FOOTNOTES

- <sup>2</sup> From solution 1 of BAILLON 75; not present in solution 2.

## Σ(1840) REFERENCES

MARTIN	77	NP	B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+)	UP
Also		NP	B126 266	B.R. Martin, M.K. Pidcock	(LOUC)	
Also		NP	B126 285	B.R. Martin, M.K. Pidcock	(LOUC)	
BAILLON	75	NP	B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL)	UP
VANHORN	75	NP	B87 145	A.J. van Horn	(LBL)	UP
Also		NP	B87 157	A.J. van Horn	(LBL)	UP
DEVENISH	74B	NP	B81 330	R.C.E. Devenish, C.D. Froggatt, B.R. Martin	(DESY+)	UP
LANGEIN	72	NP	B47 477	W. Langein, F. Wagner	(MPIM)	UP

$\Sigma(1880) \ 1/2^+$

$$I(J^P) = 1(\frac{1}{2}^+) \text{ Status: } **$$

OMITTED FROM SUMMARY TABLE

A  $P_{11}$  resonance is suggested by several partial-wave analyses, but with wide variations in the mass and other parameters. We list here all claims which lie well above the  $P_{11} \Sigma(1770)$ .

Baryon Particle Listings

$\Sigma(1880)$ ,  $\Sigma(1900)$

$\Sigma(1880)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
$\approx 1880$ OUR ESTIMATE				
1821 $\pm$ 17	ZHA NG	13A	DPWA	Multichannel
1826 $\pm$ 20	GOPAL	80	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1870 $\pm$ 10	CAMERON	78B	DPWA	$K^-p \rightarrow N\overline{K}^*$
1847 or 1863	<sup>1</sup> MARTIN	77	DPWA	$\overline{K}N$ multichannel
1960 $\pm$ 30	<sup>2</sup> BAILLON	75	IPWA	$\overline{K}N \rightarrow \Lambda\pi$
1985 $\pm$ 50	VANHORN	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
1898	<sup>3</sup> LEA	73	DPWA	Multichannel K-matrix
$\sim 1850$	ARMENTEROS	70	IPWA	$\overline{K}N \rightarrow \overline{K}N$
1950 $\pm$ 50	BARBARO-...	70	DPWA	$K^-N \rightarrow \Lambda\pi$
1920 $\pm$ 30	LITCHFIELD	70	DPWA	$K^-N \rightarrow \Lambda\pi$
1850	BAILEY	69	DPWA	$\overline{K}N \rightarrow \overline{K}N$
1882 $\pm$ 40	SMART	68	DPWA	$K^-N \rightarrow \Lambda\pi$

$\Sigma(1880)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
300± 59	ZHANG	13A	DPWA Multichannel	
86± 15	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$	
80± 10	CAMERON	78B	DPWA $K^-p \rightarrow N\overline{K}^*$	
216 or 220	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	
260± 40	<sup>2</sup> BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$	
220±140	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$	
222	<sup>3</sup> LEA	73	DPWA Multichannel K-matrix	
$\sim 30$	ARMENTEROS	70	IPWA $\overline{K}N \rightarrow \overline{K}N$	
200± 50	BARBARO-...	70	DPWA $K^-N \rightarrow \Lambda\pi$	
170± 40	LITCHFIELD	70	DPWA $K^-N \rightarrow \Lambda\pi$	
200	BAILEY	69	DPWA $\overline{K}N \rightarrow \overline{K}N$	
222±150	SMART	68	DPWA $K^-N \rightarrow \Lambda\pi$	

$\Sigma(1880)$ POLE POSITION				
REAL PART				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1776	ZHA NG	13A	DPWA	Multichannel
−2×IMAGINARY PART				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
270	ZHA NG	13A	DPWA	Multichannel

$\Sigma(1880)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\overline{K}$		
$\Gamma_2$ $\Lambda\pi$		
$\Gamma_3$ $\Sigma\pi$		
$\Gamma_4$ $\Lambda(1520)\pi$ , $D$ -wave	( 2.0±1.0) %	
$\Gamma_5$ $N\overline{K}^*(892)$ , $S=1/2$ , $P$ -wave		
$\Gamma_6$ $N\overline{K}^*(892)$ , $S=3/2$ , $P$ -wave		
$\Gamma_7$ $\Delta(1232)\overline{K}$ , $P$ -wave	(39 ±8 ) %	

$\Sigma(1880)$ BRANCHING RATIOS		
See “Sign conventions for resonance couplings” in the Note on $\Lambda$ and $\Sigma$ Resonances.		

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
0.10±0.03	ZHANG	13A	DPWA Multichannel	
0.06±0.02	GOPAL	80	DPWA $\overline{K}N \rightarrow \overline{K}N$	
0.27 or 0.27	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	
0.31	<sup>3</sup> LEA	73	DPWA Multichannel K-matrix	
0.20	ARMENTEROS	70	IPWA $\overline{K}N \rightarrow \overline{K}N$	
0.22	BAILEY	69	DPWA $\overline{K}N \rightarrow \overline{K}N$	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1880) \rightarrow \Lambda\pi$		$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.24 or −0.24	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	
−0.12 ±0.02	<sup>2</sup> BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$	
+0.05 ±0.07 −0.02	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$	
−0.169±0.119	DEVENISH	74B	Fixed- $t$ dispersion rel.	
−0.30	<sup>3</sup> LEA	73	DPWA Multichannel K-matrix	
−0.09 ±0.04	BARBARO-...	70	DPWA $K^-N \rightarrow \Lambda\pi$	
−0.14 ±0.03	LITCHFIELD	70	DPWA $K^-N \rightarrow \Lambda\pi$	
−0.11 ±0.03	SMART	68	DPWA $K^-N \rightarrow \Lambda\pi$	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1880) \rightarrow \Sigma\pi$		$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.30 or +0.29 not seen	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel	
	<sup>3</sup> LEA	73	DPWA Multichannel K-matrix	

$\Gamma(\Lambda(1520)\pi$ , $D$ -wave)/ $\Gamma_{\text{total}}$		$\Gamma_4/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.02±0.01</b>	ZHANG	13A	DPWA Multichannel	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1880) \rightarrow N\overline{K}^*(892)$ , $S=1/2$ , $P$ -wave		$(\Gamma_1\Gamma_5)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.05 ±0.03	<sup>4</sup> CAMERON	78B	DPWA $K^-p \rightarrow N\overline{K}^*$	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1880) \rightarrow N\overline{K}^*(892)$ , $S=3/2$ , $P$ -wave		$(\Gamma_1\Gamma_6)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.11 ±0.03	CAMERON	78B	DPWA $K^-p \rightarrow N\overline{K}^*$	

$\Gamma(\Delta(1232)\overline{K}$ , $P$ -wave)/ $\Gamma_{\text{total}}$		$\Gamma_7/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.39±0.08</b>	ZHANG	13A	DPWA Multichannel	

$\Sigma(1880)$ FOOTNOTES		
<sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.		
<sup>2</sup> From solution 1 of BAILLON 75; not present in solution 2.		
<sup>3</sup> Only unconstrained states from table 1 of LEA 73 are listed.		
<sup>4</sup> The published sign has been changed to be in accord with the baryon-first convention.		

$\Sigma(1880)$ REFERENCES		
ZHANG	13A	PR C88 035205
GOPAL	80	Toronto Conf. 159
CAMERON	78B	NP B146 327
MARTIN	77	NP B127 349
Also		NP B126 266
Also		NP B126 285
BAILLON	75	NP B94 39
VANHORN	75	NP B87 145
Also		NP B87 157
DEVENISH	74B	NP B81 330
LEA	73	NP B56 77
ARMENTEROS	70	Duke Conf. 123
Hyperon Resonances, 1970		
BARBARO-...	70	Duke Conf. 173
Hyperon Resonances, 1970		
LITCHFIELD	70	NP B22 269
BAILEY	69	Thesis UCRL 50617
SMART	68	PR 169 1330
H. Zhang <i>et al.</i>		(KSU)
G.P. Gopal		(RHEL) IJP
W. Cameron <i>et al.</i>		(RHEL, LOIC) IJP
B.R. Martin, M.K. Pidcock, R.G. Moorhouse		(LOUC+) IJP
B.R. Martin, M.K. Pidcock		(LOUC)
B.R. Martin, M.K. Pidcock		(LOUC) IJP
P.H. Baillon, P.J. Litchfield		(CERN, RHEL) IJP
A.J. van Horn		(LBL) IJP
A.J. van Horn		(LBL) IJP
R.C.E. Devenish, C.D. Froggatt, B.R. Martin		(DESY+) IJP
A.T. Lea <i>et al.</i>		(RHEL, LOUC, GLAS, AARH) IJP
R. Armenteros <i>et al.</i>		(CERN, HEID, SACL) IJP
A. Barbaro-Gatti		(LRL) IJP
P.J. Litchfield		(RHEL) IJP
J.M. Bailey		(LLL) IJP
W.M. Smart		(LRL) IJP

<div><math>\Sigma(1900)</math> 1/2<sup>−</sup></div>	$I(J^P) = 1(\frac{1}{2}^-)$ Status: *
OMITTED FROM SUMMARY TABLE	

$\Sigma(1900)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1900±21</b>	ZHANG	13A	DPWA Multichannel

$\Lambda(1900)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>191±47</b>	ZHANG	13A	DPWA Multichannel

$\Sigma(1900)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\overline{K}$	(67±17) %	
$\Gamma_2$ $\Sigma\pi$	(10± 5) %	

$\Sigma(1900)$ BRANCHING RATIOS				
$\Gamma(N\overline{K})/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.67±0.17</b>	ZHANG	13A	DPWA	Multichannel
$\Gamma(\Sigma\pi)/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.10±0.05</b>	ZHANG	13A	DPWA	Multichannel

$\Sigma(1900)$ REFERENCES		
ZHANG	13A	PR C88 035205
H. Zhang <i>et al.</i>		(KSU)

See key on page 601

Baryon Particle Listings  
 $\Sigma(1915), \Sigma(1940)$

$\Sigma(1915) \ 5/2^+$

$I(J^P) = 1(\frac{5}{2}^+)$  Status: \*\*\*\*

Discovered by COOL 66. For results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

This entry only includes results from partial-wave analyses. Parameters of peaks seen in cross sections and invariant-mass distributions in this region used to be listed in a separate entry immediately following. They may be found in our 1986 edition Physics Letters **170B** 1 (1986).

$\Sigma(1915)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1900 to 1935 (<math>\approx 1915</math>) OUR ESTIMATE</b>			
1920 $\pm$ 7	ZHANG 13A	DPWA	Multichannel
1937 $\pm$ 20	ALSTON-... 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
1894 $\pm$ 5	1 CORDEN 77c		$K^- n \rightarrow \Sigma \pi$
1909 $\pm$ 5	1 CORDEN 77c		$K^- n \rightarrow \Sigma \pi$
1920 $\pm$ 10	GOPAL 77	DPWA	$\bar{K}N$ multichannel
1900 $\pm$ 4	2 CORDEN 76	DPWA	$K^- n \rightarrow \Lambda \pi^-$
1920 $\pm$ 30	BAILLON 75	IPWA	$\bar{K}N \rightarrow \Lambda \pi$
1914 $\pm$ 10	HEMINGWAY 75	DPWA	$K^- p \rightarrow \bar{K}N$
1920 $\pm$ 15 -20	VANHORN 75	DPWA	$K^- p \rightarrow \Lambda \pi^0$
1920 $\pm$ 5	KANE 74	DPWA	$K^- p \rightarrow \Sigma \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
not seen	DECLAIS 77	DPWA	$\bar{K}N \rightarrow \bar{K}N$
1925 or 1933	3 MARTIN 77	DPWA	$\bar{K}N$ multichannel
1915	DEBELLEFON 76	IPWA	$K^- p \rightarrow \Lambda \pi^0$

$\Sigma(1915)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>80 to 160 (<math>\approx 120</math>) OUR ESTIMATE</b>			
149 $\pm$ 17	ZHANG 13A	DPWA	Multichannel
161 $\pm$ 20	ALSTON-... 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$
107 $\pm$ 14	1 CORDEN 77c		$K^- n \rightarrow \Sigma \pi$
85 $\pm$ 13	1 CORDEN 77c		$K^- n \rightarrow \Sigma \pi$
130 $\pm$ 10	GOPAL 77	DPWA	$\bar{K}N$ multichannel
75 $\pm$ 14	2 CORDEN 76	DPWA	$K^- n \rightarrow \Lambda \pi^-$
70 $\pm$ 20	BAILLON 75	IPWA	$\bar{K}N \rightarrow \Lambda \pi$
85 $\pm$ 15	HEMINGWAY 75	DPWA	$K^- p \rightarrow \bar{K}N$
102 $\pm$ 18	VANHORN 75	DPWA	$K^- p \rightarrow \Lambda \pi^0$
162 $\pm$ 25	KANE 74	DPWA	$K^- p \rightarrow \Sigma \pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
171 or 173	3 MARTIN 77	DPWA	$\bar{K}N$ multichannel
60	DEBELLEFON 76	IPWA	$K^- p \rightarrow \Lambda \pi^0$

$\Sigma(1915)$  POLE POSITION

<b>REAL PART</b>			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1897	ZHANG 13A	DPWA	Multichannel
<b>-2xIMAGINARY PART</b>			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
133	ZHANG 13A	DPWA	Multichannel

$\Sigma(1915)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \ N\bar{K}$	5-15 %
$\Gamma_2 \ \Lambda \pi$	seen
$\Gamma_3 \ \Sigma \pi$	seen
$\Gamma_4 \ \Sigma(1385) \pi$	<5 %
$\Gamma_5 \ \Sigma(1385) \pi, P\text{-wave}$	
$\Gamma_6 \ \Sigma(1385) \pi, F\text{-wave}$	

The above branching fractions are our estimates, not fits or averages.

$\Sigma(1915)$  BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE				
<b>0.05 to 0.15 OUR ESTIMATE</b>				
0.026 $\pm$ 0.004	ZHANG 13A	DPWA	Multichannel	
0.03 $\pm$ 0.02	4 GOPAL 80	DPWA	$\bar{K}N \rightarrow \bar{K}N$	
0.14 $\pm$ 0.05	ALSTON-... 78	DPWA	$\bar{K}N \rightarrow \bar{K}N$	
0.11 $\pm$ 0.04	HEMINGWAY 75	DPWA	$K^- p \rightarrow \bar{K}N$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.05 $\pm$ 0.03	GOPAL 77	DPWA	See GOPAL 80	
0.08 or 0.08	3 MARTIN 77	DPWA	$\bar{K}N$ multichannel	

$(\Gamma_1 \Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1915) \rightarrow \Lambda \pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1 \Gamma_2)^{1/2}/\Gamma$
VALUE				
-0.09 $\pm$ 0.03	GOPAL 77	DPWA	$\bar{K}N$ multichannel	
-0.10 $\pm$ 0.01	2 CORDEN 76	DPWA	$K^- n \rightarrow \Lambda \pi^-$	
-0.06 $\pm$ 0.02	BAILLON 75	IPWA	$\bar{K}N \rightarrow \Lambda \pi$	
-0.09 $\pm$ 0.02	VANHORN 75	DPWA	$K^- p \rightarrow \Lambda \pi^0$	
-0.087 $\pm$ 0.056	DEVENISH 74B		Fixed- $t$ dispersion rel.	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.09 or -0.09	3 MARTIN 77	DPWA	$\bar{K}N$ multichannel	
-0.10	DEBELLEFON 76	IPWA	$K^- p \rightarrow \Lambda \pi^0$	

$(\Gamma_1 \Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1915) \rightarrow \Sigma \pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1 \Gamma_3)^{1/2}/\Gamma$
VALUE				
-0.14 $\pm$ 0.01	ZHANG 13A	DPWA	Multichannel	
-0.17 $\pm$ 0.01	1 CORDEN 77c		$K^- n \rightarrow \Sigma \pi$	
-0.15 $\pm$ 0.02	1 CORDEN 77c		$K^- n \rightarrow \Sigma \pi$	
-0.19 $\pm$ 0.03	GOPAL 77	DPWA	$\bar{K}N$ multichannel	
-0.16 $\pm$ 0.03	KANE 74	DPWA	$K^- p \rightarrow \Sigma \pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.05 or -0.05	3 MARTIN 77	DPWA	$\bar{K}N$ multichannel	

$(\Gamma_1 \Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1915) \rightarrow \Sigma(1385) \pi, P\text{-wave}$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1 \Gamma_5)^{1/2}/\Gamma$
VALUE				
<0.01	CAMERON 78	DPWA	$K^- p \rightarrow \Sigma(1385) \pi$	

$(\Gamma_1 \Gamma_f)/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(1915) \rightarrow \Sigma(1385) \pi, F\text{-wave}$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1 \Gamma_6)^{1/2}/\Gamma$
VALUE				
+0.06 $\pm$ 0.02	ZHANG 13A	DPWA	Multichannel	
+0.039 $\pm$ 0.009	5 CAMERON 78	DPWA	$K^- p \rightarrow \Sigma(1385) \pi$	

$\Sigma(1915)$  FOOTNOTES

- 1 The two entries for CORDEN 77c are from two different acceptable solutions.  
2 Preferred solution 3; see CORDEN 76 for other possibilities.  
3 The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
4 The mass and width are fixed to the GOPAL 77 values due to the low elasticity.  
5 The published sign has been changed to be in accord with the baryon-first convention.

$\Sigma(1915)$  REFERENCES

ZHANG 13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
PDG 86	PL 170B 1	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)
PDG 82	PL 111B 1	M. Roos <i>et al.</i>	(HEL5, CIT, CERN)
GOPAL 80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
ALSTON-... 78	PR D18 182	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
Also	PRL 38 1007	M. Alston-Garnjost <i>et al.</i>	(LBL, MTHO+) IJP
CAMERON 78	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
CORDEN 77c	NP B125 61	M.J. Corden <i>et al.</i>	(BIRM) IJP
DECLAIS 77	CERN 77-16	Y. Declais <i>et al.</i>	(CAEN, CERN) IJP
GOPAL 77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN 77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also	NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC)
Also	NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
CORDEN 76	NP B104 382	M.J. Corden <i>et al.</i>	(BIRM) IJP
DEBELLEFON 76	NP B109 129	A. de Bellefon, A. Berthon	(CDEF) IJP
BAILLON 75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
HEMINGWAY 75	NP B91 12	R.J. Hemingway <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
VANHORN 75	NP B87 145	A.J. van Horn	(LBL) IJP
Also	NP B87 157	A.J. van Horn	(LBL) IJP
DEVENISH 74B	NP B81 330	R.C.E. Devnish, C.D. Froggatt, B.R. Martin	(DESY+) IJP
KANE 74	LBL-2452	D.F. Kane	(LBL) IJP
COOL 66	PRL 16 1228	R.L. Cool <i>et al.</i>	(BNL)

$\Sigma(1940) \ 3/2^+$

$I(J^P) = 1(\frac{3}{2}^+)$  Status: \*

OMITTED FROM SUMMARY TABLE

$\Sigma(1940)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1941 <math>\pm</math> 18</b>	ZHANG 13A	DPWA	Multichannel

# Baryon Particle Listings

## $\Sigma(1940)$ , $\Sigma(1940)$

### $\Lambda(1945)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>400\pm49</math></b>	ZHANG	13A	DPWA Multichannel

### $\Sigma(1940)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\overline{K}$	( $13.0\pm2.0$ ) %
$\Gamma_2$ $\Sigma\pi$	( $4.0\pm2.0$ ) %
$\Gamma_3$ $\Sigma(1385)\pi$ , $P$ -wave	( $22\pm7$ ) %
$\Gamma_4$ $\Lambda(1520)\pi$ , $S$ -wave	( $5.0\pm2.0$ ) %

### $\Sigma(1940)$ BRANCHING RATIOS

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.13\pm0.02</math></b>	ZHA NG	13A	DPWA Multichannel

$\Gamma(\Sigma\pi)/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.04\pm0.02</math></b>	ZHANG	13A	DPWA Multichannel

$\Gamma(\Sigma(1385)\pi, P\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.22\pm0.07</math></b>	ZHA NG	13A	DPWA Multichannel

$\Gamma(\Lambda(1520)\pi, S\text{-wave})/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.05 \pm 0.02</math></b>	ZHA NG	13A	DPWA Multichannel

### $\Sigma(1940)$ REFERENCES

ZHANG	13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
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$$\Sigma(1940) \ 3/2^-$$

$$I(J^P) = 1(\frac{3}{2}^-) \text{ Status: } ***$$

For results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

Not all analyses require this state. It is not required by the GOYAL 77 analysis of  $K^-n \rightarrow (\Sigma\pi)^-$  nor by the GOPAL 80 analysis of  $K^-n \rightarrow K^-n$ . See also HEMINGWAY 75.

### $\Sigma(1940)$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1900 to 1950 (<math>\approx 1940</math>) OUR ESTIMATE</b>			
$1920\pm50$	GOPAL	77	DPWA $\overline{K}N$ multichannel
$1950\pm30$	BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$
$1949^{+40}_{-60}$	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
$1935\pm80$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
$1940\pm20$	LITCHFIELD	74B	DPWA $K^-p \rightarrow \Lambda(1520)\pi^0$
$1950\pm20$	LITCHFIELD	74C	DPWA $K^-p \rightarrow \Delta(1232)\overline{K}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1886 or 1893	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel
1940	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0, F_{17}$ wave

### $\Sigma(1940)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>150 to 300 (<math>\approx 220</math>) OUR ESTIMATE</b>			
$170\pm25$	CAMERON	78B	DPWA $K^-p \rightarrow N\overline{K}^*$
$300\pm80$	GOPAL	77	DPWA $\overline{K}N$ multichannel
$150\pm75$	BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$
$160^{+70}_{-40}$	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
$330\pm80$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
$60\pm20$	LITCHFIELD	74B	DPWA $K^-p \rightarrow \Lambda(1520)\pi^0$
$70^{+30}_{-20}$	LITCHFIELD	74C	DPWA $K^-p \rightarrow \Delta(1232)\overline{K}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
157 or 159	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

### $\Sigma(1940)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $N\overline{K}$	<20 %
$\Gamma_2$ $\Lambda\pi$	seen
$\Gamma_3$ $\Sigma\pi$	seen
$\Gamma_4$ $\Sigma(1385)\pi$	seen
$\Gamma_5$ $\Sigma(1385)\pi$ , $S$ -wave	
$\Gamma_6$ $\Lambda(1520)\pi$	seen
$\Gamma_7$ $\Lambda(1520)\pi$ , $P$ -wave	
$\Gamma_8$ $\Lambda(1520)\pi$ , $F$ -wave	
$\Gamma_9$ $\Delta(1232)\overline{K}$	seen
$\Gamma_{10}$ $\Delta(1232)\overline{K}$ , $S$ -wave	
$\Gamma_{11}$ $\Delta(1232)\overline{K}$ , $D$ -wave	
$\Gamma_{12}$ $N\overline{K}^*(892)$	seen
$\Gamma_{13}$ $N\overline{K}^*(892)$ , $S=3/2$ , $S$ -wave	

### $\Sigma(1940)$ BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.2 OUR ESTIMATE</b>			
<0.04	GOPAL	77	DPWA $\overline{K}N$ multichannel
0.14 or 0.13	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Lambda\pi$	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.06 \pm 0.03$	GOPAL	77	DPWA $\overline{K}N$ multichannel
$-0.04 \pm 0.02$	BAILLON	75	IPWA $\overline{K}N \rightarrow \Lambda\pi$
$-0.05 \pm^{+0.03}_{-0.02}$	VANHORN	75	DPWA $K^-p \rightarrow \Lambda\pi^0$
$-0.153 \pm 0.070$	DEVENISH	74B	Fixed- $t$ dispersion rel.
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.15$ or $-0.14$	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

$(\Gamma_1\Gamma_3)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.08 \pm 0.04$	GOPAL	77	DPWA $\overline{K}N$ multichannel
$-0.14 \pm 0.04$	KANE	74	DPWA $K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$+0.16$ or $+0.16$	<sup>1</sup> MARTIN	77	DPWA $\overline{K}N$ multichannel

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Lambda(1520)\pi$ , <i>P-wave</i>	$(\Gamma_1\Gamma_7)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$< 0.03$	CAMERON	77	DPWA $K^-p \rightarrow \Lambda(1520)\pi^0$
$-0.11 \pm 0.04$	LITCHFIELD	74B	DPWA $K^-p \rightarrow \Lambda(1520)\pi^0$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Lambda(1520)\pi, F\text{-wave}$				$(\Gamma_1\Gamma_8)^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.062 \pm 0.021$	CAMERON	77	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$
$-0.08 \pm 0.04$	LITCHFIELD	74B	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$

$(\Gamma_1\Gamma_9)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Delta(1232)\overline{K}, S\text{-wave}$	$(\Gamma_1\Gamma_{10})^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$-0.16 \pm 0.05$	LITCHFIELD	74C	DPWA $K^-p \rightarrow \Delta(1232)\overline{K}$

$(\Gamma_1\Gamma_7)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Delta(1232)\overline{K}, D\text{-wave}$				$(\Gamma_1\Gamma_{11})^{1/2}/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
$-0.14 \pm 0.05$	LITCHFIELD	74C	DPWA	$K^-p \rightarrow \Delta(1232)\overline{K}$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow \Sigma(1385)\pi$	$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$+0.066\pm0.025$	<sup>2</sup> CAMERON	78	DPWA $K^-p \rightarrow \Sigma(1385)\pi$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(1940) \rightarrow N\overline{K}^*(892)$				$(\Gamma_1\Gamma_{12})^{1/2}/\Gamma$
VALUE	DOCUMENT ID		TECN	COMMENT
$-0.09 \pm 0.02$	<sup>3</sup> CAMERON	78B	DPWA	$K^-p \rightarrow N\overline{K}^*$

### $\Sigma(1940)$ FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> The published sign has been changed to be in accord with the baryon-first convention.  
<sup>3</sup> Upper limits on the  $D_1$  and  $D_3$  waves are each 0.03.

See key on page 601

Baryon Particle Listings  
 $\Sigma(1940)$ ,  $\Sigma(2000)$ ,  $\Sigma(2030)$

$\Sigma(1940)$  REFERENCES

PDG	82	PL 111B 1	M. Roos <i>et al.</i>	(HEL5, CIT, CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal	(RHEL)
CAMERON	78	NP B143 189	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
CAMERON	78B	NP B146 327	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
CAMERON	77	NP B131 399	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
GOYAL	77	PR D16 2746	D.P. Goyal, A.V. Sodhi	(DELH)
MARTIN	77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also		NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC)
Also		NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
DEBELLEFON	76	NP B109 129	A. de Bellefon, A. Berthon	(CDEF) IJP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
HEMINGWAY	75	NP B91 12	R.J. Hemingway <i>et al.</i>	(CERN, HEIDH, MPIM) IJP
VANHORN	75	NP B87 145	A.J. van Horn	(LBL) IJP
Also		NP B87 157	A.J. van Horn	(LBL) IJP
DEVENISH	74B	NP B81 330	R.C.E. Devenish, C.D. Froggatt, B.R. Martin	(DESY+) IJP
KANE	74	LBL-2452	D.F. Kane	(LBL) IJP
LITCHFIELD	74B	NP B74 19	P.J. Litchfield <i>et al.</i>	(CERN, HEIDH) IJP
LITCHFIELD	74C	NP B74 39	P.J. Litchfield <i>et al.</i>	(CERN, HEIDH) IJP

$\Sigma(2000) 1/2^-$

$I(J^P) = 1(\frac{1}{2}^-)$  Status: \*

OMITTED FROM SUMMARY TABLE

We list here all reported  $S_{11}$  states lying above the  $\Sigma(1750) S_{11}$ .  
ZHANG 13A finds no evidence for those states.

$\Sigma(2000)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 2000$ OUR ESTIMATE			
1944 $\pm$ 15	GOPAL 80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
1955 $\pm$ 15	GOPAL 77	DPWA	$\bar{K}N$ multichannel
1755 or 1834	<sup>1</sup> MARTIN 77	DPWA	$\bar{K}N$ multichannel
2004 $\pm$ 40	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda\pi^0$

$\Sigma(2000)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
215 $\pm$ 25	GOPAL 80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
170 $\pm$ 40	GOPAL 77	DPWA	$\bar{K}N$ multichannel
413 or 450	<sup>1</sup> MARTIN 77	DPWA	$\bar{K}N$ multichannel
116 $\pm$ 40	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda\pi^0$

$\Sigma(2000)$  DECAY MODES

Mode
$\Gamma_1$ $N\bar{K}$
$\Gamma_2$ $\Lambda\pi$
$\Gamma_3$ $\Sigma\pi$
$\Gamma_4$ $\Lambda(1520)\pi$
$\Gamma_5$ $N\bar{K}^*(892)$ , $S=1/2$ , $S$ -wave
$\Gamma_6$ $N\bar{K}^*(892)$ , $S=3/2$ , $D$ -wave

$\Sigma(2000)$  BRANCHING RATIOS

See "Sign conventions for resonance couplings" in the Note on  $\Lambda$  and  $\Sigma$  Resonances.

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE				
0.51 $\pm$ 0.05	GOPAL 80	DPWA	$\bar{K}N \rightarrow \bar{K}N$	
0.44 $\pm$ 0.05	GOPAL 77	DPWA	See GOPAL 80	
0.62 or 0.57	<sup>1</sup> MARTIN 77	DPWA	$\bar{K}N$ multichannel	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2000) \rightarrow \Lambda\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$
VALUE				
0.08 $\pm$ 0.03	GOPAL 77	DPWA	$\bar{K}N$ multichannel	
-0.19 or -0.18	<sup>1</sup> MARTIN 77	DPWA	$\bar{K}N$ multichannel	
not seen	BAILLON 75	IPWA	$\bar{K}N \rightarrow \Lambda\pi$	
+0.07 $\pm$ 0.02	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda\pi^0$	
-0.01				

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2000) \rightarrow \Sigma\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$
VALUE				
+0.20 $\pm$ 0.04	GOPAL 77	DPWA	$\bar{K}N$ multichannel	
+0.26 or +0.24	<sup>1</sup> MARTIN 77	DPWA	$\bar{K}N$ multichannel	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2000) \rightarrow \Lambda(1520)\pi$	DOCUMENT ID	TECN	COMMENT	$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$
VALUE				
+0.081 $\pm$ 0.021	<sup>2</sup> CAMERON 77	DPWA	$P$ -wave decay	

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2000) \rightarrow N\bar{K}^*(892)$ , $S=1/2$ , $S$ -wave $(\Gamma_1\Gamma_5)^{1/2}/\Gamma$	DOCUMENT ID	TECN	COMMENT
VALUE			
+0.10 $\pm$ 0.02	<sup>2</sup> CAMERON 78B	DPWA	$K^-p \rightarrow N\bar{K}^*$

$(\Gamma_1\Gamma_f)^{1/2}/\Gamma_{\text{total}}$  in  $N\bar{K} \rightarrow \Sigma(2000) \rightarrow N\bar{K}^*(892)$ ,  $S=3/2$ ,  $D$ -wave

$(\Gamma_1\Gamma_6)^{1/2}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
-0.07 $\pm$ 0.03	CAMERON 78B	DPWA	$K^-p \rightarrow N\bar{K}^*$

$\Sigma(2000)$  FOOTNOTES

- <sup>1</sup> The two MARTIN 77 values are from a T-matrix pole and from a Breit-Wigner fit.  
<sup>2</sup> The published sign has been changed to be in accord with the baryon-first convention.

$\Sigma(2000)$  REFERENCES

ZHANG 13A	PR C88 035205	H. Zhang <i>et al.</i>	(KSU)
GOPAL 80	Toronto Conf. 159	G.P. Gopal	(RHEL) IJP
CAMERON 78B	NP B146 327	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
CAMERON 77	NP B131 399	W. Cameron <i>et al.</i>	(RHEL, LOIC) IJP
GOPAL 77	NP B119 362	G.P. Gopal <i>et al.</i>	(LOIC, RHEL) IJP
MARTIN 77	NP B127 349	B.R. Martin, M.K. Pidcock, R.G. Moorhouse	(LOUC+) IJP
Also	NP B126 266	B.R. Martin, M.K. Pidcock	(LOUC)
Also	NP B126 285	B.R. Martin, M.K. Pidcock	(LOUC) IJP
BAILLON 75	NP B94 39	P.H. Baillon, P.J. Litchfield	(CERN, RHEL) IJP
VANHORN 75	NP B87 145	A.J. van Horn	(LBL) IJP
Also	NP B87 157	A.J. van Horn	(LBL) IJP

$\Sigma(2030) 7/2^+$

$I(J^P) = 1(\frac{7}{2}^+)$  Status: \*\*\*

Discovered by COOL 66 and by WOHL 66. For most results published before 1974 (they are now obsolete), see our 1982 edition Physics Letters **111B** 1 (1982).

This entry only includes results from partial-wave analyses. Parameters of peaks seen in cross sections and invariant-mass distributions around 2030 MeV may be found in our 1984 edition, Reviews of Modern Physics **56** S1 (1984).

$\Sigma(2030)$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2025 to 2040 (<math>\approx 2030</math>) OUR ESTIMATE</b>			
2030 $\pm$ 5	ZHANG 13A	DPWA	Multichannel
2036 $\pm$ 5	GOPAL 80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
2038 $\pm$ 10	CORDEN 77B		$K^-N \rightarrow N\bar{K}^*$
2040 $\pm$ 5	GOPAL 77	DPWA	$\bar{K}N$ multichannel
2030 $\pm$ 3	<sup>1</sup> CORDEN 76	DPWA	$K^-n \rightarrow \Lambda\pi^-$
2035 $\pm$ 15	BAILLON 75	IPWA	$\bar{K}N \rightarrow \Lambda\pi$
2038 $\pm$ 10	HEMINGWAY 75	DPWA	$K^-p \rightarrow \bar{K}N$
2042 $\pm$ 11	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
2020 $\pm$ 6	KANE 74	DPWA	$K^-p \rightarrow \Sigma\pi$
2035 $\pm$ 10	LITCHFIELD 74B	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$
2020 $\pm$ 30	LITCHFIELD 74C	DPWA	$K^-p \rightarrow \Delta(1232)\bar{K}$
2025 $\pm$ 10	LITCHFIELD 74D	DPWA	$K^-p \rightarrow \Lambda(1820)\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2027 to 2057	GOYAL 77	DPWA	$K^-N \rightarrow \Sigma\pi$
2030	DEBELLEFON 76	IPWA	$K^-p \rightarrow \Lambda\pi^0$

$\Sigma(2030)$  WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>150 to 200 (<math>\approx 180</math>) OUR ESTIMATE</b>			
207 $\pm$ 17	ZHANG 13A	DPWA	Multichannel
172 $\pm$ 10	GOPAL 80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
137 $\pm$ 40	CORDEN 77B		$K^-N \rightarrow N\bar{K}^*$
190 $\pm$ 10	GOPAL 77	DPWA	$\bar{K}N$ multichannel
201 $\pm$ 9	<sup>1</sup> CORDEN 76	DPWA	$K^-n \rightarrow \Lambda\pi^-$
180 $\pm$ 20	BAILLON 75	IPWA	$\bar{K}N \rightarrow \Lambda\pi$
172 $\pm$ 15	HEMINGWAY 75	DPWA	$K^-p \rightarrow \bar{K}N$
178 $\pm$ 13	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
111 $\pm$ 5	KANE 74	DPWA	$K^-p \rightarrow \Sigma\pi$
160 $\pm$ 20	LITCHFIELD 74B	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$
200 $\pm$ 30	LITCHFIELD 74C	DPWA	$K^-p \rightarrow \Delta(1232)\bar{K}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
260	DECLAIS 77	DPWA	$\bar{K}N \rightarrow \bar{K}N$
126 to 195	GOYAL 77	DPWA	$K^-N \rightarrow \Sigma\pi$
160	DEBELLEFON 76	IPWA	$K^-p \rightarrow \Lambda\pi^0$
70 to 125	LITCHFIELD 74D	DPWA	$K^-p \rightarrow \Lambda(1820)\pi^0$

$\Sigma(2030)$  POLE POSITION

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1993	ZHANG 13A	DPWA	Multichannel

# Baryon Particle Listings

## $\Sigma(2030)$ , $\Sigma(2070)$

−2×IMAGINARY PART			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
176	ZHANG	13A	DPWA Multichannel

$\Sigma(2030)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$ $N\bar{K}$	17–23 %	
$\Gamma_2$ $\Lambda\pi$	17–23 %	
$\Gamma_3$ $\Sigma\pi$	5–10 %	
$\Gamma_4$ $\Xi K$	<2 %	
$\Gamma_5$ $\Sigma(1385)\pi$	5–15 %	
$\Gamma_6$ $\Sigma(1385)\pi$ , $F$ -wave		
$\Gamma_7$ $\Lambda(1520)\pi$	10–20 %	
$\Gamma_8$ $\Lambda(1520)\pi$ , $D$ -wave		
$\Gamma_9$ $\Lambda(1520)\pi$ , $G$ -wave		
$\Gamma_{10}$ $\Delta(1232)\bar{K}$	10–20 %	
$\Gamma_{11}$ $\Delta(1232)\bar{K}$ , $F$ -wave		
$\Gamma_{12}$ $\Delta(1232)\bar{K}$ , $H$ -wave		
$\Gamma_{13}$ $N\bar{K}^*(892)$	<5 %	
$\Gamma_{14}$ $N\bar{K}^*(892)$ , $S=1/2$ , $F$ -wave		
$\Gamma_{15}$ $N\bar{K}^*(892)$ , $S=3/2$ , $F$ -wave		
$\Gamma_{16}$ $\Lambda(1820)\pi$ , $P$ -wave		
The above branching fractions are our estimates, not fits or averages.		

$\Sigma(2030)$ BRANCHING RATIOS	
See “Sign conventions for resonance couplings” in the Note on $\Lambda$ and $\Sigma$ Resonances.	

$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.17 to 0.23 OUR ESTIMATE				
0.13±0.01	ZHANG	13A	DPWA	Multichannel
0.19±0.03	GOPAL	80	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.18±0.03	HEMINGWAY	75	DPWA	$K^-p \rightarrow \bar{K}N$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.15	DECLAIS	77	DPWA	$\bar{K}N \rightarrow \bar{K}N$
0.24±0.02	GOPAL	77	DPWA	See GOPAL 80

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Lambda\pi$	$(\Gamma_1\Gamma_2)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.15 ±0.01	ZHANG	13A	DPWA	Multichannel
+0.18 ±0.02	GOPAL	77	DPWA	$\bar{K}N$ multichannel
+0.20 ±0.01	<sup>1</sup> CORDEN	76	DPWA	$K^-n \rightarrow \Lambda\pi^-$
+0.18 ±0.02	BAILLON	75	IPWA	$\bar{K}N \rightarrow \Lambda\pi$
+0.20 ±0.01	VANHORN	75	DPWA	$K^-p \rightarrow \Lambda\pi^0$
+0.195±0.053	DEVENISH	74B	Fixed- $t$ dispersion rel.	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.20	DEBELLEFON	76	IPWA	$K^-p \rightarrow \Lambda\pi^0$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_3)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
−0.08 ±0.01	ZHANG	13A	DPWA	Multichannel
−0.09 ±0.01	<sup>2</sup> CORDEN	77C	$K^-n \rightarrow \Sigma\pi$	
−0.06 ±0.01	<sup>2</sup> CORDEN	77C	$K^-n \rightarrow \Sigma\pi$	
−0.15 ±0.03	GOPAL	77	DPWA	$\bar{K}N$ multichannel
−0.10 ±0.01	KANE	74	DPWA	$K^-p \rightarrow \Sigma\pi$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
−0.085±0.02	<sup>3</sup> GOYAL	77	DPWA	$K^-N \rightarrow \Sigma\pi$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Xi K$	$(\Gamma_1\Gamma_4)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.023	MULLER	69B	DPWA	$K^-p \rightarrow \Xi K$
<0.05	BURGUN	68	DPWA	$K^-p \rightarrow \Xi K$
<0.05	TRIPP	67	RVUE	$K^-p \rightarrow \Xi K$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Sigma(1385)\pi$ , $F$ -wave	$(\Gamma_1\Gamma_6)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.16 ±0.01	ZHANG	13A	DPWA	Multichannel
+0.153±0.026	<sup>4</sup> CAMERON	78	DPWA	$K^-p \rightarrow \Sigma(1385)\pi$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Lambda(1520)\pi$ , $D$ -wave	$(\Gamma_1\Gamma_8)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.114±0.010	<sup>4</sup> CAMERON	77	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$
0.14 ±0.03	LITCHFIELD	74B	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.10 ±0.03	<sup>5</sup> CORDEN	75B	DBC	$K^-n \rightarrow N\bar{K}\pi^-$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Lambda(1520)\pi$ , $G$ -wave	$(\Gamma_1\Gamma_9)^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.146±0.010	<sup>4</sup> CAMERON	77	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$
0.02 ±0.02	LITCHFIELD	74B	DPWA	$K^-p \rightarrow \Lambda(1520)\pi^0$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Delta(1232)\bar{K}$ , $F$ -wave	$(\Gamma_1\Gamma_{11})^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.12±0.02	ZHANG	13A	DPWA	Multichannel
0.16±0.03	LITCHFIELD	74C	DPWA	$K^-p \rightarrow \Delta(1232)\bar{K}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.17±0.03	<sup>5</sup> CORDEN	75B	DBC	$K^-n \rightarrow N\bar{K}\pi^-$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Delta(1232)\bar{K}$ , $H$ -wave	$(\Gamma_1\Gamma_{12})^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.00±0.02	LITCHFIELD	74C	DPWA	$K^-p \rightarrow \Delta(1232)\bar{K}$

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow N\bar{K}^*(892)$ , $S=1/2$ , $F$ -wave	$(\Gamma_1\Gamma_{14})^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.06±0.02	ZHANG	13A	DPWA	Multichannel
+0.06±0.03	<sup>4</sup> CAMERON	78B	DPWA	$K^-p \rightarrow N\bar{K}^*$
−0.02±0.01	CORDEN	77B	$K^-d \rightarrow NN\bar{K}^*$	

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow N\bar{K}^*(892)$ , $S=3/2$ , $F$ -wave	$(\Gamma_1\Gamma_{15})^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
+0.05±0.01	ZHANG	13A	DPWA	Multichannel
+0.04±0.03	<sup>6</sup> CAMERON	78B	DPWA	$K^-p \rightarrow N\bar{K}^*$
−0.12±0.02	CORDEN	77B	$K^-d \rightarrow NN\bar{K}^*$	

$(\Gamma_i\Gamma_f)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2030) \rightarrow \Lambda(1820)\pi$ , $P$ -wave	$(\Gamma_1\Gamma_{16})^{1/2}/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.14±0.02	CORDEN	75B	DBC	$K^-n \rightarrow N\bar{K}\pi^-$
0.18±0.04	LITCHFIELD	74D	DPWA	$K^-p \rightarrow \Lambda(1820)\pi^0$

$\Sigma(2030)$ FOOTNOTES	
<sup>1</sup> Preferred solution 3; see CORDEN 76 for other possibilities.	
<sup>2</sup> The two entries for CORDEN 77C are from two different acceptable solutions.	
<sup>3</sup> This coupling is extracted from unnormalized data.	
<sup>4</sup> The published sign has been changed to be in accord with the baryon-first convention.	
<sup>5</sup> An upper limit.	
<sup>6</sup> The upper limit on the $G_3$ wave is 0.03.	

$\Sigma(2030)$ REFERENCES			
ZHANG	13A	PR C08 035205	H. Zhang <i>et al.</i> (KSU)
PDG	84	RMP 56 51	C. G. Wohl <i>et al.</i> (LBL, CIT. CERN)
PDG	82	PL 111B 1	M. Roos <i>et al.</i> (HELS, CIT. CERN)
GOPAL	80	Toronto Conf. 159	G.P. Gopal (RHEL)JUP
CAMERON	78	NP B143 189	W. Cameron <i>et al.</i> (RHEL, LOIC)JUP
CAMERON	78B	NP B146 327	W. Cameron <i>et al.</i> (RHEL, LOIC)JUP
CAMERON	77	NP B131 399	W. Cameron <i>et al.</i> (RHEL, LOIC)JUP
CORDEN	77B	NP B121 365	M.J. Corden <i>et al.</i> (BIRM)JUP
CORDEN	77C	NP B125 61	M.J. Corden <i>et al.</i> (BIRM)JUP
DECLAIS	77	CERN 77-16	Y. Declais <i>et al.</i> (CAEN, CERN)JUP
GOPAL	77	NP B119 362	G.P. Gopal <i>et al.</i> (LOIC, RHEL)JUP
GOYAL	77	PR D16 2746	D.P. Goyal, A.V. Sodhi (DELH)JUP
CORDEN	76	NP B104 382	M.J. Corden <i>et al.</i> (BIRM)JUP
DEBELLEFON	76	NP B109 129	A. de Bellefon, A. Berthon (CDEF)JUP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield (CERN, RHEL)JUP
CORDEN	75B	NP B92 365	M.J. Corden <i>et al.</i> (BIRM)JUP
HEMINGWAY	75	NP B91 12	R.J. Hemingway <i>et al.</i> (CERN, HEIDH, MPIM)JUP
VANHORN	75	NP B87 145	A.J. van Horn (LBL)JUP
Also		NP B87 157	A.J. van Horn (LBL)JUP
DEVENISH	74B	NP B81 330	R.C.E. Devenish, C.D. Froggatt, B.R. Martin (DES+)JUP
KANE	74	LBL-2452	D.F. Kane (LBL)JUP
LITCHFIELD	74B	NP B74 19	P.J. Litchfield <i>et al.</i> (CERN, HEIDH)JUP
LITCHFIELD	74C	NP B74 39	P.J. Litchfield <i>et al.</i> (CERN, HEIDH)JUP
LITCHFIELD	74D	NP B74 12	P.J. Litchfield <i>et al.</i> (CERN, HEIDH)JUP
MULLER	69B	Thesis UCRL 19372	R.A. Muller (LRL)JUP
BURGUN	68	NP B8 447	G. Burgun <i>et al.</i> (SACL, CDEF, RHEL)JUP
TRIPP	67	NP B3 10	R.D. Tripp <i>et al.</i> (LRL, SLAC, CERN+)JUP
COOL	66	PRL 16 1228	R.L. Cool <i>et al.</i> (BNL)JUP
WOHL	66	PRL 17 107	C.G. Wohl, F.T. Solmitz, M.L. Stevenson (LRL)JUP

$\Sigma(2070) 5/2^+$

$I(J^P) = 1(\frac{5}{2}^+)$  Status: \*

OMITTED FROM SUMMARY TABLE

This state suggested by BERTHON 70B finds support in GOPAL 80 with new  $K^-p$  polarization and  $K^-n$  angular distributions. The very broad state seen in KANE 72 is not required in the later (KANE 74) analysis of  $\bar{K}N \rightarrow \Sigma\pi$ .



See key on page 601

Baryon Particle Listings

$\Sigma(2070)$ ,  $\Sigma(2080)$ ,  $\Sigma(2100)$ ,  $\Sigma(2250)$

$\Sigma(2070)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 2070$ OUR ESTIMATE			
2051 $\pm$ 25	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
2057	KANE	72	DPWA $K^-p \rightarrow \Sigma\pi$
2070 $\pm$ 10	BERTHON	70B	DPWA $K^-p \rightarrow \Sigma\pi$

$\Sigma(2070)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
300 $\pm$ 30	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
906	KANE	72	DPWA $K^-p \rightarrow \Sigma\pi$
140 $\pm$ 20	BERTHON	70B	DPWA $K^-p \rightarrow \Sigma\pi$

$\Sigma(2070)$ DECAY MODES	
Mode	
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Sigma\pi$

$\Sigma(2070)$ BRANCHING RATIOS			
See "Sign conventions for resonance couplings" in the Note on $\Lambda$ and $\Sigma$ Resonances.			
$\Gamma(N\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
0.08 $\pm$ 0.03	GOPAL	80	DPWA $\bar{K}N \rightarrow \bar{K}N$
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2070) \rightarrow \Sigma\pi$	DOCUMENT ID	TECN	COMMENT
+0.104	KANE	72	DPWA $K^-p \rightarrow \Sigma\pi$
+0.12 $\pm$ 0.02	BERTHON	70B	DPWA $K^-p \rightarrow \Sigma\pi$

$\Sigma(2070)$ REFERENCES			
GOPAL	80	Toronto Conf. 159	G.P. Gopal (RHEL) IJP
KANE	74	LBL-2452	D.F. Kane (LBL) IJP
KANE	72	PR D5 1583	D.F.J. Kane (LBL) IJP
BERTHON	70B	NP B24 417	A. Berthon et al. (CDEF, RHEL, SACL) IJP

$\Sigma(2080)$ 3/2 <sup>+</sup>	$I(J^P) = 1(\frac{3}{2}^+)$ Status: **
OMITTED FROM SUMMARY TABLE	
Suggested by some but not all partial-wave analyses across this region.	

$\Sigma(2080)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 2080$ OUR ESTIMATE			
2091 $\pm$ 7	1 CORDEN	76	DPWA $K^-n \rightarrow \Lambda\pi^-$
2070 to 2120	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$
2120 $\pm$ 40	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$ (sol. 1)
2140 $\pm$ 40	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$ (sol. 2)
2082 $\pm$ 4	COX	70	DPWA See CORDEN 76
2070 $\pm$ 30	LITCHFIELD	70	DPWA $K^-N \rightarrow \Lambda\pi$

$\Sigma(2080)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
186 $\pm$ 48	1 CORDEN	76	DPWA $K^-n \rightarrow \Lambda\pi^-$
100	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$
240 $\pm$ 50	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$ (sol. 1)
200 $\pm$ 50	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$ (sol. 2)
87 $\pm$ 20	COX	70	DPWA See CORDEN 76
250 $\pm$ 40	LITCHFIELD	70	DPWA $K^-N \rightarrow \Lambda\pi$

$\Sigma(2080)$ DECAY MODES	
Mode	
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Lambda\pi$

$\Sigma(2080)$ BRANCHING RATIOS			
See "Sign conventions for resonance couplings" in the Note on $\Lambda$ and $\Sigma$ Resonances.			
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2080) \rightarrow \Lambda\pi$	DOCUMENT ID	TECN	COMMENT
-0.10 $\pm$ 0.03	1 CORDEN	76	DPWA $K^-n \rightarrow \Lambda\pi^-$
-0.10	DEBELLEFON	76	IPWA $K^-p \rightarrow \Lambda\pi^0$
-0.13 $\pm$ 0.04	BAILLON	75	IPWA $\bar{K}N \rightarrow \Lambda\pi$ (sol. 1 and 2)
-0.16 $\pm$ 0.03	COX	70	DPWA See CORDEN 76
-0.09 $\pm$ 0.03	LITCHFIELD	70	DPWA $K^-N \rightarrow \Lambda\pi$

$\Sigma(2080)$ FOOTNOTES	
1 Preferred solution 3; see CORDEN 76 for other possibilities, including a $D_{15}$ at this mass.	

$\Sigma(2080)$ REFERENCES			
CORDEN	76	NP B104 382	M.J. Corden et al. (BIRM) IJP
DEBELLEFON	76	NP B109 129	A. de Bellefon, A. Berthon (CDEF) IJP
		Also NP B90 1	A. de Bellefon et al. (CDEF, SACL) IJP
BAILLON	75	NP B94 39	P.H. Baillon, P.J. Litchfield (CERN, RHEL) IJP
COX	70	NP B19 61	G.F. Cox et al. (BIRM, EDIN, GLAS, LOIC) IJP
LITCHFIELD	70	NP B22 269	P.J. Litchfield (RHEL) IJP

$\Sigma(2100)$ 7/2 <sup>-</sup>	$I(J^P) = 1(\frac{7}{2}^-)$ Status: *
OMITTED FROM SUMMARY TABLE	

$\Sigma(2100)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$\approx 2100$ OUR ESTIMATE			
2060 $\pm$ 20	BARBARO...	70	DPWA $K^-p \rightarrow \Lambda\pi^0$
2120 $\pm$ 30	BARBARO...	70	DPWA $K^-p \rightarrow \Sigma\pi$

$\Sigma(2100)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
70 $\pm$ 30	BARBARO...	70	DPWA $K^-p \rightarrow \Lambda\pi^0$
135 $\pm$ 30	BARBARO...	70	DPWA $K^-p \rightarrow \Sigma\pi$

$\Sigma(2100)$ DECAY MODES	
Mode	
$\Gamma_1$	$N\bar{K}$
$\Gamma_2$	$\Lambda\pi$
$\Gamma_3$	$\Sigma\pi$

$\Sigma(2100)$ BRANCHING RATIOS			
See "Sign conventions for resonance couplings" in the Note on $\Lambda$ and $\Sigma$ Resonances.			
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2100) \rightarrow \Lambda\pi$	DOCUMENT ID	TECN	COMMENT
-0.07 $\pm$ 0.02	BARBARO...	70	DPWA $K^-p \rightarrow \Lambda\pi^0$
$(\Gamma_1\Gamma_2)^{1/2}/\Gamma_{\text{total}}$ in $N\bar{K} \rightarrow \Sigma(2100) \rightarrow \Sigma\pi$	DOCUMENT ID	TECN	COMMENT
+0.13 $\pm$ 0.02	BARBARO...	70	DPWA $K^-p \rightarrow \Sigma\pi$

$\Sigma(2100)$ REFERENCES			
BARBARO...	70	Duke Conf. 173	A. Barbaro-Galtieri (LRL) IJP
Hyperon Resonances, 1970			

$\Sigma(2250)$	$I(J^P) = 1(?^?)$ Status: ***
Results from partial-wave analyses are too weak to warrant separating them from the production and cross-section experiments. LASINSKI 71 in $\bar{K}N$ using a Pomeron + resonances model, and DEBELLEFON 76, DEBELLEFON 77, and DEBELLEFON 78 in energy-dependent partial-wave analyses of $\bar{K}N \rightarrow \Lambda\pi$ , $\Sigma\pi$ , and $N\bar{K}$ , respectively, suggest two resonances around this mass.	

# Baryon Particle Listings

## $\Sigma(2250)$ , $\Sigma(2455)$ Bumps

$\Sigma(2250)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2210 to 2280 (<math>\approx 2250</math>) OUR ESTIMATE</b>			
2270 $\pm$ 50	DEBELLEFON 78	DPWA	$D_5$ wave
2210 $\pm$ 30	DEBELLEFON 78	DPWA	$G_9$ wave
2275 $\pm$ 20	DEBELLEFON 77	DPWA	$D_5$ wave
2215 $\pm$ 20	DEBELLEFON 77	DPWA	$G_9$ wave
2300 $\pm$ 30	<sup>1</sup> DEBELLEFON 75B	HBC	$K^-p \rightarrow \Xi^*0 K^0$
2251 $\pm$ <sub>-20</sub> <sup>+30</sup>	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda \pi^0, F_5$ wave
2280 $\pm$ 14	AGUILAR-... 70B	HBC	$K^-p$ 3.9, 4.6 GeV/ $c$
2237 $\pm$ 11	BRICMAN 70	CNTR	Total, charge exchange
2255 $\pm$ 10	COOL 70	CNTR	$K^-p, K^-d$ total
2250 $\pm$ 7	BUGG 68	CNTR	$K^-p, K^-d$ total
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2260	DEBELLEFON 76	IPWA	$D_5$ wave
2215	DEBELLEFON 76	IPWA	$G_9$ wave
2250 $\pm$ 20	LU 70	CNTR	$\gamma p \rightarrow K^+ Y^*$
2245	BLANPIED 65	CNTR	$\gamma p \rightarrow K^+ Y^*$
2299 $\pm$ 6	BOCK 65	HBC	$\overline{p}p$ 5.7 GeV/ $c$

$\Sigma(2250)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>60 to 150 (<math>\approx 100</math>) OUR ESTIMATE</b>			
120 $\pm$ 40	DEBELLEFON 78	DPWA	$D_5$ wave
80 $\pm$ 20	DEBELLEFON 78	DPWA	$G_9$ wave
70 $\pm$ 20	DEBELLEFON 77	DPWA	$D_5$ wave
60 $\pm$ 20	DEBELLEFON 77	DPWA	$G_9$ wave
130 $\pm$ 20	<sup>1</sup> DEBELLEFON 75B	HBC	$K^-p \rightarrow \Xi^*0 K^0$
192 $\pm$ 30	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda \pi^0, F_5$ wave
100 $\pm$ 20	AGUILAR-... 70B	HBC	$K^-p$ 3.9, 4.6 GeV/ $c$
164 $\pm$ 50	BRICMAN 70	CNTR	Total, charge exchange
230 $\pm$ 20	BUGG 68	CNTR	$K^-p, K^-d$ total
• • • We do not use the following data for averages, fits, limits, etc. • • •			
100	DEBELLEFON 76	IPWA	$D_5$ wave
140	DEBELLEFON 76	IPWA	$G_9$ wave
170	COOL 70	CNTR	$K^-p, K^-d$ total
125	LU 70	CNTR	$\gamma p \rightarrow K^+ Y^*$
150	BLANPIED 65	CNTR	$\gamma p \rightarrow K^+ Y^*$
21 $\pm$ <sub>-21</sub> <sup>+17</sup>	BOCK 65	HBC	$\overline{p}p$ 5.7 GeV/ $c$

$\Sigma(2250)$ DECAY MODES		
Mode	Fraction ( $\Gamma_i/\Gamma$ )	
$\Gamma_1$	$N\overline{K}$	<10 %
$\Gamma_2$	$\Lambda\pi$	seen
$\Gamma_3$	$\Sigma\pi$	seen
$\Gamma_4$	$N\overline{K}\pi$	
$\Gamma_5$	$\Xi(1530)K$	
The above branching fractions are our estimates, not fits or averages.		

$\Sigma(2250)$ BRANCHING RATIOS	
See "Sign conventions for resonance couplings" in the Note on $\Lambda$ and $\Sigma$ Resonances.	

$\Gamma(N\overline{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.1 OUR ESTIMATE</b>			
0.08 $\pm$ 0.02	DEBELLEFON 78	DPWA	$D_5$ wave
0.02 $\pm$ 0.01	DEBELLEFON 78	DPWA	$G_9$ wave

$(J+\frac{1}{2})\times\Gamma(N\overline{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.16 $\pm$ 0.12	BRICMAN 70	CNTR	Total, charge exchange
0.42	COOL 70	CNTR	$K^-p, K^-d$ total
0.47	BUGG 68	CNTR	

$(\Gamma_1\Gamma_f)^{\frac{1}{2}}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(2250) \rightarrow \Lambda\pi$	$(\Gamma_1\Gamma_2)^{\frac{1}{2}}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
-0.16 $\pm$ 0.03	VANHORN 75	DPWA	$K^-p \rightarrow \Lambda \pi^0, F_5$ wave
• • • We do not use the following data for averages, fits, limits, etc. • • •			
+0.11	DEBELLEFON 76	IPWA	$D_5$ wave
-0.10	DEBELLEFON 76	IPWA	$G_9$ wave
-0.18	BARBARO-... 70	DPWA	$K^-p \rightarrow \Lambda \pi^0, G_9$ wave

$(\Gamma_1\Gamma_f)^{\frac{1}{2}}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(2250) \rightarrow \Sigma\pi$	$(\Gamma_1\Gamma_3)^{\frac{1}{2}}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
+0.06 $\pm$ 0.02	DEBELLEFON 77	DPWA	$D_5$ wave
-0.03 $\pm$ 0.02	DEBELLEFON 77	DPWA	$G_9$ wave
+0.07	BARBARO-... 70	DPWA	$K^-p \rightarrow \Sigma\pi, G_9$ wave

$\Gamma(N\overline{K})/\Gamma(\Sigma\pi)$	$\Gamma_1/\Gamma_3$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.18	BARNES 69	HBC	1 standard dev. limit

$\Gamma(\Lambda\pi)/\Gamma(\Sigma\pi)$	$\Gamma_2/\Gamma_3$		
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.18	BARNES 69	HBC	1 standard dev. limit

$(\Gamma_1\Gamma_f)^{\frac{1}{2}}/\Gamma_{\text{total}}$ in $N\overline{K} \rightarrow \Sigma(2250) \rightarrow \Xi(1530)K$	$(\Gamma_1\Gamma_5)^{\frac{1}{2}}/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
0.18 $\pm$ 0.04	<sup>1</sup> DEBELLEFON 75B	HBC	$K^-p \rightarrow \Xi^*0 K^0$

$\Sigma(2250)$ FOOTNOTES	
<sup>1</sup> Seen in the (initial and final state) $D_5$ wave. Isospin not determined.	

$\Sigma(2250)$ REFERENCES			
DEBELLEFON 78	NC 42A 403	A. de Bellefon <i>et al.</i>	(CDEF, SACL)JJP
DEBELLEFON 77	NC 37A 175	A. de Bellefon <i>et al.</i>	(CDEF, SACL)JJP
DEBELLEFON 76	NP B109 129	A. de Bellefon, A. Berthon	(CDEF)JJP
Also	NP B30 1	A. de Bellefon <i>et al.</i>	(CDEF, SACL)JJP
DEBELLEFON 75B	NC 28A 289	A. de Bellefon <i>et al.</i>	(CDEF, SACL)JJP
VANHORN 75	NP B87 145	A.J. van Horn	(LBL)JJP
Also	NP B87 157	A.J. van Horn	(LBL)JJP
LASINSKI 71	NP B29 125	T.A. Lasinski	(EFI)JJP
AGUILAR-... 70B	PRL 25 58	M. Aguilar-Benitez <i>et al.</i>	(BNL, SYR)A
BARBARO-... 70	Duke Conf. 173	A. Barbaro-Gallieri	(LRL)JJP
Hyperon Resonances, 1970			
BRICMAN 70	PL 31B 152	C. Bricman <i>et al.</i>	(CERN, CAEN, SACL)
COOL 70	PR D1 1887	R.L. Cool <i>et al.</i>	(BNL)I
Also	PRL 16 1228	R.L. Cool <i>et al.</i>	(BNL)I
LU 70	PR D2 1846	D.C. Lu <i>et al.</i>	(YALE)
BARNES 69	PRL 22 479	V.E. Barnes <i>et al.</i>	(BNL, SYR)A
BUGG 68	PR 168 1466	D.V. Bugg <i>et al.</i>	(RHEL, BIRM, CAVE)I
BLANPIED 65	PRL 14 741	W.A. Blanpied <i>et al.</i>	(YALE, CEA)
BOCK 65	PL 17 166	R.K. Bock <i>et al.</i>	(CERN, SACL)

$\Sigma(2455)$ Bumps	$I(J^P) = 1(?^?)$	Status: **
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OMITTED FROM SUMMARY TABLE	
There is also some slight evidence for $Y^*$ states in this mass region from the reaction $\gamma p \rightarrow K^+ X$ — see GREENBERG 68.	

$\Sigma(2455)$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 2455</math> OUR ESTIMATE</b>			
2455 $\pm$ 10	ABRAMS 70	CNTR	$K^-p, K^-d$ total
2455 $\pm$ 7	BUGG 68	CNTR	$K^-p, K^-d$ total

$\Sigma(2455)$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
140	ABRAMS 70	CNTR	$K^-p, K^-d$ total
100 $\pm$ 20	BUGG 68	CNTR	

$\Sigma(2455)$ DECAY MODES	
Mode	
$\Gamma_1$ $N\overline{K}$	

$\Sigma(2455)$ BRANCHING RATIOS				
$(J+\frac{1}{2})\times\Gamma(N\overline{K})/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
0.39	ABRAMS	70	CNTR	$K^-p, K^-d$ total
$0.05\pm0.05$	<sup>1</sup> BRICMAN	70	CNTR	Total, charge exchange
0.3	BUGG	68	CNTR	

$\Sigma(2455)$ FOOTNOTES	
<sup>1</sup> Fit of total cross section given by BRICMAN 70 is poor in this region.	

$\Sigma(2455)$ REFERENCES			
ABRAMS 70	PR D1 1917	R.J. Abrams <i>et al.</i>	(BNL)I
Also	PRL 19 478	R.J. Abrams <i>et al.</i>	(BNL)
BRICMAN 70	PL 31B 152	C. Bricman <i>et al.</i>	(CERN, CAEN, SACL)
BUGG 68	PR 168 1466	D.V. Bugg <i>et al.</i>	(RHEL, BIRM, CAVE)I
GREENBERG 68	PRL 20 221	J.S. Greenberg <i>et al.</i>	(YALE)

See key on page 601

Baryon Particle Listings

$\Sigma(2620)$  Bumps,  $\Sigma(3000)$  Bumps,  $\Sigma(3170)$  Bumps

$\Sigma(2620)$  Bumps

$I(J^P) = 1(?^?)$  Status: \*\*

OMITTED FROM SUMMARY TABLE

$\Sigma(2620)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
$\approx 2620$ OUR ESTIMATE				
2542±22	DIBIANCA	75	DBC	$K^- N \rightarrow \Xi K \pi$
2620±15	ABRAMS	70	CNTR	$K^- p, K^- d$ total

$\Sigma(2620)$ WIDTH				
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT	
221±81	DIBIANCA	75	DBC	$K^- N \rightarrow \Xi K \pi$
175	ABRAMS	70	CNTR	$K^- p, K^- d$ total

$\Sigma(2620)$ DECAY MODES	
Mode	
$\Gamma_1$	$N \bar{K}$

$\Sigma(2620)$ BRANCHING RATIOS				
$(J+\frac{1}{2}) \times \Gamma(N \bar{K}) / \Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
0.32	ABRAMS	70	CNTR	$K^- p, K^- d$ total
0.36±0.12	BRICMAN	70	CNTR	Total, charge exchange

$\Sigma(2620)$ REFERENCES				
DIBIANCA	75	NP B98 137	F.A. Dibianca, R.J. Endorf	(CMU)
ABRAMS	70	PR D1 1917	R.J. Abrams <i>et al.</i>	(BNL)1
Also		PRL 19 678	R.J. Abrams <i>et al.</i>	(BNL)
BRICMAN	70	PL 31B 152	C. Bricman <i>et al.</i>	(CERN, CAEN, SACL)

$\Sigma(3000)$  Bumps

$I(J^P) = 1(?^?)$  Status: \*

OMITTED FROM SUMMARY TABLE

Seen as an enhancement in  $\Lambda \pi$  and  $\bar{K} N$  invariant mass spectra and in the missing mass of neutrals recoiling against a  $K^0$ .

$\Sigma(3000)$ MASS				
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
$\approx 3000$ OUR ESTIMATE				
3000	EHRlich	66	HBC	0 $\pi^- p$ 7.91 GeV/c

$\Sigma(3000)$ DECAY MODES	
Mode	
$\Gamma_1$	$N \bar{K}$
$\Gamma_2$	$\Lambda \pi$

$\Sigma(3000)$ REFERENCES				
EHRlich	66	PR 152 1194	R. Ehrlich, W. Selove, H. Yuta	(PENN)1

$\Sigma(3170)$  Bumps

$I(J^P) = 1(?^?)$  Status: \*

OMITTED FROM SUMMARY TABLE

Seen by AMIRZADEH 79 as a narrow 6.5-standard-deviation enhancement in the reaction  $K^- p \rightarrow Y^{*+} \pi^-$  using data from independent high statistics bubble chamber experiments at 8.25 and 6.5 GeV/c. The dominant decay modes are multibody, multistrange final states and the production is via isospin-3/2 baryon exchange. Isospin 1 is favored.

Not seen in a  $K^- p$  experiment in LASS at 11 GeV/c (ASTON 85B).

$\Sigma(3170)$ MASS (PRODUCTION EXPERIMENTS)				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$\approx 3170$ OUR ESTIMATE				
3170±5	35	AMIRZADEH 79	HBC	$K^- p \rightarrow Y^{*+} \pi^-$

$\Sigma(3170)$ WIDTH (PRODUCTION EXPERIMENTS)				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<20	35	1 AMIRZADEH 79	HBC	$K^- p \rightarrow Y^{*+} \pi^-$

$\Sigma(3170)$ DECAY MODES (PRODUCTION EXPERIMENTS)	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$	$\Lambda K \bar{K} \pi$ 's seen
$\Gamma_2$	$\Sigma K \bar{K} \pi$ 's seen
$\Gamma_3$	$\Xi K \pi$ 's seen

$\Sigma(3170)$ BRANCHING RATIOS (PRODUCTION EXPERIMENTS)				
$\Gamma(\Lambda K \bar{K} \pi \text{'s}) / \Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
VALUE				
seen	AMIRZADEH 79	HBC	$K^- p \rightarrow Y^{*+} \pi^-$	
$\Gamma(\Sigma K \bar{K} \pi \text{'s}) / \Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
VALUE				
seen	AMIRZADEH 79	HBC	$K^- p \rightarrow Y^{*+} \pi^-$	
$\Gamma(\Xi K \pi \text{'s}) / \Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	
VALUE				
seen	AMIRZADEH 79	HBC	$K^- p \rightarrow Y^{*+} \pi^-$	

$\Sigma(3170)$ FOOTNOTES (PRODUCTION EXPERIMENTS)	
1 Observed width consistent with experimental resolution.	

$\Sigma(3170)$ REFERENCES (PRODUCTION EXPERIMENTS)				
ASTON	85B	PR D32 2270	D. Aston <i>et al.</i>	(SLAC, CARL, CNRC, CINC)
AMIRZADEH	79	PL 89B 125	J. Amirzadeh <i>et al.</i>	(BIRM, CERN, GLAS+)1
Also		Toronto Conf. 263	J.B. Kinson <i>et al.</i>	(BIRM, CERN, GLAS+)1

Baryon Particle Listings

$\Xi^0$

$\Xi$  BARYONS  
( $S = -2$ ,  $I = 1/2$ )  
 $\Xi^0 = uss$ ,  $\Xi^- = dss$

$\Xi^0$

$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \*\*\*\*

The parity has not actually been measured, but + is of course expected.

$\Xi^0$  MASS

The fit uses the  $\Xi^0$ ,  $\Xi^-$ , and  $\Xi^+$  masses and the  $\Xi^- - \Xi^0$  mass difference. It assumes that the  $\Xi^-$  and  $\Xi^+$  masses are the same.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1314.86±0.20 OUR FIT</b>				
<b>1314.82±0.06±0.20</b>	3120	FANTI	00	NA48 $p$ Be, 450 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1315.2 ±0.92	49	WILQUET	72	HLBC
1313.4 ±1.8	1	PALMER	68	HBC

$m_{\Xi^-} - m_{\Xi^0}$

The fit uses the  $\Xi^0$ ,  $\Xi^-$ , and  $\Xi^+$  masses and the  $\Xi^- - \Xi^0$  mass difference. It assumes that the  $\Xi^-$  and  $\Xi^+$  masses are the same.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>6.85±0.21 OUR FIT</b>				
<b>6.3 ±0.7 OUR AVERAGE</b>				
6.9 ±2.2	29	LONDON	66	HBC
6.1 ±0.9	88	PJERROU	65B	HBC
6.8 ±1.6	23	JAUNEAU	63	FBC
• • • We do not use the following data for averages, fits, limits, etc. • • •				
6.1 ±1.6	45	CARMONY	64B	HBC See PJERROU 65B

$\Xi^0$  MEAN LIFE

VALUE ( $10^{-10}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.90±0.09 OUR AVERAGE</b>				
2.83±0.16	6300	<sup>1</sup> ZECH	77	SPEC Neutral hyperon beam
$2.86^{+0.21}_{-0.19}$	652	BALTAY	74	HBC 1.75 GeV/c $K^- p$
$2.90^{+0.32}_{-0.27}$	157	<sup>2</sup> MAYEUR	72	HLBC 2.1 GeV/c $K^-$
$3.07^{+0.22}_{-0.20}$	340	DAUBER	69	HBC
3.0 ±0.5	80	PJERROU	65B	HBC
$2.5^{+0.4}_{-0.3}$	101	HUBBARD	64	HBC
$3.9^{+1.4}_{-0.8}$	24	JAUNEAU	63	FBC
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$3.5^{+1.0}_{-0.8}$	45	CARMONY	64B	HBC See PJERROU 65B
<sup>1</sup> The ZECH 77 result is $\tau_{\Xi^0} = [2.77 - (\tau_A - 2.69)] \times 10^{-10}$ s, in which we use $\tau_A = 2.63 \times 10^{-10}$ s.				
<sup>2</sup> The MAYEUR 72 value is modified by the erratum.				

$\Xi^0$  MAGNETIC MOMENT

See the “Note on Baryon Magnetic Moments” in the  $\Lambda$  Listings.

VALUE ( $\mu_N$ )	EVTS	DOCUMENT ID	TECN
<b>-1.250±0.014 OUR AVERAGE</b>			
-1.253±0.014	270k	COX	81 SPEC
-1.20 ±0.06	42k	BUNCE	79 SPEC

$\Xi^0$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \Lambda\pi^0$	(99.524±0.012) %	
$\Gamma_2 \Lambda\gamma$	( 1.17 ±0.07 ) $\times 10^{-3}$	
$\Gamma_3 \Lambda e^+ e^-$	( 7.6 ±0.6 ) $\times 10^{-6}$	
$\Gamma_4 \Sigma^0 \gamma$	( 3.33 ±0.10 ) $\times 10^{-3}$	
$\Gamma_5 \Sigma^+ e^- \bar{\nu}_e$	( 2.52 ±0.08 ) $\times 10^{-4}$	
$\Gamma_6 \Sigma^+ \mu^- \bar{\nu}_\mu$	( 2.33 ±0.35 ) $\times 10^{-6}$	

$\Delta S = \Delta Q$  (SQ) violating modes or  
 $\Delta S = 2$  forbidden (S2) modes

$\Gamma_7 \Sigma^- e^+ \nu_e$	SQ	< 9	$\times 10^{-4}$	90%
$\Gamma_8 \Sigma^- \mu^+ \nu_\mu$	SQ	< 9	$\times 10^{-4}$	90%
$\Gamma_9 p \pi^-$	S2	< 8	$\times 10^{-6}$	90%
$\Gamma_{10} p e^- \bar{\nu}_e$	S2	< 1.3	$\times 10^{-3}$	
$\Gamma_{11} p \mu^- \bar{\nu}_\mu$	S2	< 1.3	$\times 10^{-3}$	

CONSTRAINED FIT INFORMATION

An overall fit to 5 branching ratios uses 11 measurements and one constraint to determine 5 parameters. The overall fit has a  $\chi^2 = 7.5$  for 7 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	−57			
$x_4$	−82	0		
$x_5$	−7	0	0	
$x_6$	0	0	0	1
	$x_1$	$x_2$	$x_4$	$x_5$

$\Xi^0$  BRANCHING RATIOS

$\Gamma(\Lambda\gamma)/\Gamma(\Lambda\pi^0)$				$\Gamma_2/\Gamma_1$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.17±0.07 OUR FIT</b>				
<b>1.17±0.07 OUR AVERAGE</b>				
1.17±0.05±0.06	672	<sup>3</sup> LAI	04A	NA48 p Be, 450 GeV
1.91±0.34±0.19	31	<sup>4</sup> FANTI	00	NA48 p Be, 450 GeV
1.06±0.12±0.11	116	JAMES	90	SPEC FNAL hyperons
<sup>3</sup> LAI 04A used our 2002 value of 99.5% for the $\Xi^0 \rightarrow \Lambda\pi^0$ branching fraction to get $\Gamma(\Xi^0 \rightarrow \Lambda\gamma)/\Gamma_{\text{total}} = (1.16 \pm 0.05 \pm 0.06) \times 10^{-3}$ . We adjust slightly to go back to what was directly measured.				
<sup>4</sup> FANTI 00 used our 1998 value of 99.5% for the $\Xi^0 \rightarrow \Lambda\pi^0$ branching fraction to get $\Gamma(\Xi^0 \rightarrow \Lambda\gamma)/\Gamma_{\text{total}} = (1.90 \pm 0.34 \pm 0.19) \times 10^{-3}$ . We adjust slightly to go back to what was directly measured.				

$\Gamma(\Lambda e^+ e^-)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.6±0.4±0.5</b>	397 ± 21	<sup>5</sup> BATLEY	07c	NA48 p Be, 400 GeV
<sup>5</sup> This BATLEY 07c result is consistent with internal bremsstrahlung.				

$\Gamma(\Sigma^0\gamma)/\Gamma(\Lambda\pi^0)$				$\Gamma_4/\Gamma_1$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.35±0.10 OUR FIT</b>				
<b>3.35±0.10 OUR AVERAGE</b>				
3.34±0.05±0.09	4045	ALAVI-HARATI	01c	KTEV $p$ nucleus, 800 GeV
3.16±0.76±0.32	17	<sup>6</sup> FANTI	00	NA48 $p$ Be, 450 GeV
3.56±0.42±0.10	85	TEIGE	89	SPEC FNAL hyperons
<sup>6</sup> FANTI 00 used our 1998 value of 99.5% for the $\Xi^0 \rightarrow \Lambda\pi^0$ branching fraction to get $\Gamma(\Xi^0 \rightarrow \Sigma^0\gamma)/\Gamma_{\text{total}} = (3.14 \pm 0.76 \pm 0.32) \times 10^{-3}$ . We adjust slightly to go back to what was directly measured.				

$\Gamma(\Sigma^+ e^- \bar{\nu}_e)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.52±0.08 OUR FIT</b>				
<b>2.53±0.08 OUR AVERAGE</b>				
2.51±0.03±0.09	6101	BATLEY	07	NA48 $p$ Be, 400 GeV
2.55±0.14±0.10	419	<sup>7</sup> BATLEY	07	NA48 $p$ Be, 400 GeV
2.71±0.22±0.31	176	AFFOLDER	99	KTEV $p$ nucleus, 800 GeV
<sup>7</sup> This BATLEY 07 result is for $\Xi^0 \rightarrow \Sigma^- e^+ \nu_e$ events.				

$\Gamma(\Sigma^+ \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$	$\Gamma_6/\Gamma$			
VALUE (units $10^{-6}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.3 ± 0.4 OUR FIT</b>				
<b>2.17 ± 0.32 ± 0.17</b>	66	<sup>8</sup> BATLEY	13	NA48 p Be, 400 GeV
<sup>8</sup> BATLEY 13 used $\Xi^0 \rightarrow \Sigma^+ e^- \bar{\nu}_e$ decay as a normalization mode and its branching fraction value of $(2.51 \pm 0.03 \pm 0.09) \times 10^{-4}$ from BATLEY 07.				

$\Gamma(\Sigma^+ \mu^- \bar{\nu}_\mu)/\Gamma(\Sigma^+ e^- \bar{\nu}_e)$					$\Gamma_6/\Gamma_5$
VALUE	EVTS		DOCUMENT ID	TECN	COMMENT
0.0092±0.0015 OUR FIT					
0.018 $\begin{smallmatrix} +0.007 \\ -0.005 \end{smallmatrix}$ ±0.002	9	ABOUZAID	05	KTEV	p nucleus 800 GeV

$\Gamma(\Sigma^- e^+ \nu_e)/\Gamma(\Lambda\pi^0)$ Test of $\Delta S = \Delta Q$ rule.					$\Gamma_7/\Gamma_1$	
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<0.9	90	0	YEH	74	HBC	Effective denom.=2500
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<1.5			DAUBER	69	HBC	
<6			HUBBARD	66	HBC	

$\Gamma(\Sigma^- \mu^+ \nu_\mu)/\Gamma(\Lambda\pi^0)$ Test of $\Delta S = \Delta Q$ rule.					$\Gamma_8/\Gamma_1$	
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<0.9	90	0	YEH	74	HBC	Effective denom.=2500
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<1.5			DAUBER	69	HBC	
<6			HUBBARD	66	HBC	

$\Gamma(p\pi^-)/\Gamma(\Lambda\pi^0)$ $\Delta S=2$ . Forbidden in first-order weak interaction.					$\Gamma_9/\Gamma_1$	
VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
< 8.2	90		WHITE	05	HYCP	p Cu, 800 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •						
< 36	90		GEWENIGER	75	SPEC	
<1800	90	0	YEH	74	HBC	Effective denom.=1300
< 900			DAUBER	69	HBC	
<5000			HUBBARD	66	HBC	

$\Gamma(p e^- \bar{\nu}_e)/\Gamma(\Lambda\pi^0)$ $\Delta S=2$ . Forbidden in first-order weak interaction.					$\Gamma_{10}/\Gamma_1$	
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<1.3			DAUBER	69	HBC	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<3.4	90	0	YEH	74	HBC	Effective denom.=670
<6			HUBBARD	66	HBC	

$\Gamma(p\mu^- \bar{\nu}_\mu)/\Gamma(\Lambda\pi^0)$ $\Delta S=2$ . Forbidden in first-order weak interaction.					$\Gamma_{11}/\Gamma_1$	
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<1.3			DAUBER	69	HBC	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<3.5	90	0	YEH	74	HBC	Effective denom.=664
<6			HUBBARD	66	HBC	

 **$\Xi^0$  DECAY PARAMETERS**

See the "Note on Baryon Decay Parameters" in the neutron Listings.

$\alpha(\Xi^0) \alpha_{-}(\Lambda)$ This is a product of the $\Xi^0 \rightarrow \Lambda\pi^0$ and $\Lambda \rightarrow p\pi^-$ asymmetries.						
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>-0.261 ± 0.006 OUR AVERAGE</b>						
-0.276 ± 0.001 ± 0.035	4M	BATLEY	10B	NA48	p Be, 400 GeV	
-0.260 ± 0.004 ± 0.005	300k	HANDLER	82	SPEC	FNAL hyperons	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
-0.317 ± 0.027	6075	BUNCE	78	SPEC	FNAL hyperons	
-0.35 ± 0.06	505	BALTAY	74	HBC	$K^- p$ 1.75 GeV/c	
-0.28 ± 0.06	739	DAUBER	69	HBC	$K^- p$ 1.7-2.6 GeV/c	

$\alpha$ FOR $\Xi^0 \rightarrow \Lambda\pi^0$ The above average, $\alpha(\Xi^0)\alpha_{-}(\Lambda) = -0.261 \pm 0.006$ , divided by our current average $\alpha_{-}(\Lambda) = 0.642 \pm 0.013$ , gives the following value for $\alpha(\Xi^0)$ .						
VALUE	DOCUMENT ID					
<b>-0.406 ± 0.013 OUR EVALUATION</b>						

$\phi$ ANGLE FOR $\Xi^0 \rightarrow \Lambda\pi^0$ (tan $\phi = \beta/\gamma$ )				
VALUE (°)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>21 ± 12 OUR AVERAGE</b>				
16 ± 17	652	BALTAY	74	HBC 1.75 GeV/c $K^- p$
38 ± 19	739	DAUBER	69	HBC
- 8 ± 30	146	BERGE	66	HBC

<sup>9</sup> DAUBER 69 uses  $\alpha_{\Lambda} = 0.647 \pm 0.020$ .  
<sup>10</sup> The errors have been multiplied by 1.2 due to approximations used for the  $\Xi$  polarization; see DAUBER 69 for a discussion.

**RADIATIVE HYPERON DECAYS**

Revised July 2011 by J.D. Jackson (LBNL).

The weak radiative decays of spin-1/2 hyperons,  $B_i \rightarrow B_f \gamma$ , yield information about matrix elements (form factors) similar to that gained from weak hadronic decays. For a polarized spin-1/2 hyperon decaying radiatively via a  $\Delta Q = 0$ ,  $\Delta S = 1$

transition, the angular distribution of the direction  $\hat{\mathbf{p}}$  of the final spin-1/2 baryon in the hyperon rest frame is

$$\frac{dN}{d\Omega} = \frac{N}{4\pi} (1 + \alpha_{\gamma} \mathbf{P}_i \cdot \hat{\mathbf{p}}) . \quad (1)$$

Here  $\mathbf{P}_i$  is the polarization of the decaying hyperon, and  $\alpha_{\gamma}$  is the asymmetry parameter. In terms of the form factors  $F_1(q^2)$ ,  $F_2(q^2)$ , and  $G(q^2)$  of the effective hadronic weak electromagnetic vertex,

$$F_1(q^2)\gamma_{\lambda} + iF_2(q^2)\sigma_{\lambda\mu}q^{\mu} + G(q^2)\gamma_{\lambda}\gamma_5 ,$$

$\alpha_{\gamma}$  is

$$\alpha_{\gamma} = \frac{2 \operatorname{Re}[G(0)F_M^*(0)]}{|G(0)|^2 + |F_M(0)|^2} , \quad (2)$$

where  $F_M = (m_i - m_f)[F_2 - F_1/(m_i + m_f)]$ . If the decaying hyperon is unpolarized, the decay baryon has a longitudinal polarization given by  $P_f = -\alpha_{\gamma}$  [1].

The angular distribution for the weak hadronic decay,  $B_i \rightarrow B_f \pi$ , has the same form as Eq. (1), but of course with a different asymmetry parameter,  $\alpha_{\pi}$ . Now, however, if the decaying hyperon is unpolarized, the decay baryon has a longitudinal polarization given by  $P_f = +\alpha_{\pi}$  [2,3]. The difference of sign is because the spins of the pion and photon are different.

**$\Xi^0 \rightarrow \Lambda \gamma$  decay**—The radiative decay  $\Xi^0 \rightarrow \Lambda \gamma$  of an unpolarized  $\Xi^0$  uses the hadronic decay  $\Lambda \rightarrow p\pi^-$  as the analyzer. As noted above, the longitudinal polarization of the  $\Lambda$  will be  $P_{\Lambda} = -\alpha_{\Xi\Lambda\gamma}$ . Let  $\alpha_{-}$  be the  $\Lambda \rightarrow p\pi^-$  asymmetry parameter and  $\theta_{\Lambda p}$  be the angle, as seen in the  $\Lambda$  rest frame, between the  $\Lambda$  line of flight and the proton momentum. Then the hadronic version of Eq. (1) applied to the  $\Lambda \rightarrow p\pi^-$  decay gives

$$\frac{dN}{d\cos\theta_{\Lambda p}} = \frac{N}{2} (1 - \alpha_{\Xi\Lambda\gamma} \alpha_{-} \cos\theta_{\Lambda p}) \quad (3)$$

for the angular distribution of the proton in the  $\Lambda$  frame. Our current value, from the CERN NA48/1 experiment [4], is  $\alpha_{\Xi\Lambda\gamma} = -0.704 \pm 0.019 \pm 0.064$ .

**$\Xi^0 \rightarrow \Sigma^0 \gamma$  decay**—The asymmetry parameter here,  $\alpha_{\Xi\Sigma\gamma}$ , is measured by following the decay chain  $\Xi^0 \rightarrow \Sigma^0 \gamma$ ,  $\Sigma^0 \rightarrow \Lambda \gamma$ ,  $\Lambda \rightarrow p\pi^-$ . Again, for an unpolarized  $\Xi^0$ , the longitudinal polarization of the  $\Sigma^0$  will be  $P_{\Sigma} = -\alpha_{\Xi\Sigma\gamma}$ . In the  $\Sigma^0 \rightarrow \Lambda \gamma$  decay, a parity-conserving magnetic-dipole transition, the polarization of the  $\Sigma^0$  is transferred to the  $\Lambda$ , as may be seen as follows. Let  $\theta_{\Sigma\Lambda}$  be the angle seen in the  $\Sigma^0$  rest frame between the  $\Sigma^0$  line of flight and the  $\Lambda$  momentum. For  $\Sigma^0$  helicity +1/2, the probability amplitudes for positive and negative spin states of the  $\Sigma^0$  along the  $\Lambda$  momentum are  $\cos(\theta_{\Sigma\Lambda}/2)$  and  $\sin(\theta_{\Sigma\Lambda}/2)$ . Then the amplitude for a negative helicity photon and a negative helicity  $\Lambda$  is  $\cos(\theta_{\Sigma\Lambda}/2)$ , while the amplitude for positive helicities for the photon and  $\Lambda$  is  $\sin(\theta_{\Sigma\Lambda}/2)$ . For  $\Sigma^0$  helicity -1/2, the amplitudes are interchanged. If the  $\Sigma^0$  has longitudinal polarization  $P_{\Sigma}$ , the probabilities for  $\Lambda$  helicities  $\pm 1/2$  are therefore

$$p(\pm 1/2) = \frac{1}{2}(1 \mp P_{\Sigma}) \cos^2(\theta_{\Sigma\Lambda}/2) + \frac{1}{2}(1 \pm P_{\Sigma}) \sin^2(\theta_{\Sigma\Lambda}/2) , \quad (4)$$

# Baryon Particle Listings

 $\Xi^0, \Xi^-$ 

and the longitudinal polarization of the  $\Lambda$  is

$$P_{\Lambda} = -P_{\Sigma} \cos \theta_{\Sigma\Lambda} = +\alpha_{\Xi\Sigma\gamma} \cos \theta_{\Sigma\Lambda} . \quad (5)$$

Using Eq. (1) for the  $\Lambda \rightarrow p\pi^-$  decay again, we get for the joint angular distribution of the  $\Sigma^0 \rightarrow \Lambda\gamma$ ,  $\Lambda \rightarrow p\pi^-$  chain,

$$\frac{d^2N}{d\cos\theta_{\Sigma\Lambda}d\cos\theta_{\Lambda p}} = \frac{N}{4} (1 + \alpha_{\Xi\Sigma\gamma} \cos\theta_{\Sigma\Lambda} \alpha_- \cos\theta_{\Lambda p}) . \quad (6)$$

Our current average for  $\alpha_{\Xi\Sigma\gamma}$  is  $-0.69 \pm 0.06$  [4,5].

## References

1. R.E. Behrends, Phys. Rev. **111**, 1691 (1958); see Eq. (7) or (8).
2. In ancient times, the signs of the asymmetry term in the angular distributions of radiative and hadronic decays of polarized hyperons were sometimes opposite. For roughly 50 years, however, the overwhelming convention has been to make them the same. The aim, not always achieved, is to remove ambiguities.
3. For the definition of  $\alpha_{\pi}$ , see the note on “Baryon Decay Parameters” in the Neutron Listings.
4. J.R. Batley *et al.*, Phys. Lett. **B693**, 241 (2010).
5. A. Alavi-Harati *et al.*, Phys. Rev. Lett. **86**, 3239 (2001).

## $\alpha$ FOR $\Xi^0 \rightarrow \Lambda\gamma$

See the note above on “Radiative Hyperon Decays.”

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.704 \pm 0.019 \pm 0.064</math></b>	52k	<sup>11</sup> BATLEY	10b NA48	$p$ Be, 400 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$-0.78 \pm 0.18 \pm 0.06$	672	LAI	04A NA48	See BATLEY 10b
$-0.43 \pm 0.44$	87	<sup>12</sup> JAMES	90 SPEC	FNAL hyperons

<sup>11</sup> BATLEY 10b also measured the  $\Xi^0 \rightarrow \bar{\Lambda}\gamma$  asymmetry to be  $-0.798 \pm 0.064$  (no systematic error given) with 4769 events.

<sup>12</sup> The sign has been changed; see the erratum, JAMES 02.

## $\alpha$ FOR $\Xi^0 \rightarrow \Lambda e^+ e^-$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.8 \pm 0.2</math></b>	397 $\pm$ 21	<sup>13</sup> BATLEY	07c NA48	$p$ Be, 400 GeV

<sup>13</sup> This BATLEY 07c result is consistent with the asymmetry  $\alpha$  for  $\Xi^0 \rightarrow \Lambda\gamma$ , as expected if the mechanism is internal bremsstrahlung.

## $\alpha$ FOR $\Xi^0 \rightarrow \Sigma^0\gamma$

See the note above on “Radiative Hyperon Decays.”

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.69 \pm 0.06</math> OUR AVERAGE</b>				
$-0.729 \pm 0.030 \pm 0.076$	15k	<sup>14</sup> BATLEY	10b NA48	$p$ Be, 400 GeV
$-0.63 \pm 0.08 \pm 0.05$	4045	ALAVI-HARATI01c	KTEV	$p$ nucleus, 800 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$+0.20 \pm 0.32 \pm 0.05$	85	<sup>15</sup> TEIGE	89 SPEC	FNAL hyperons

<sup>14</sup> BATLEY 10b also measured the  $\Xi^0 \rightarrow \bar{\Sigma}^0\gamma$  asymmetry to be  $-0.786 \pm 0.104$  (no systematic error given) with 1404 events.

<sup>15</sup> This result has been withdrawn, due to an error. See the erratum, TEIGE 02.

## $g_2(0)/f_1(0)$ FOR $\Xi^0 \rightarrow \Sigma^+ e^- \bar{\nu}_e$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.22 \pm 0.05</math> OUR AVERAGE</b>				
$1.21 \pm 0.05$		BATLEY	13 NA48	$p$ Be, 400 GeV
$1.32^{+0.21}_{-0.17} \pm 0.05$	487	<sup>16</sup> ALAVI-HARATI01i	KTEV	$p$ nucleus, 800 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1.20 \pm 0.04 \pm 0.03$	6520	<sup>17</sup> BATLEY	07 NA48	See BATLEY 13

<sup>16</sup> ALAVI-HARATI 01i assumes here that the second-class current is zero and that the weak-magnetism term takes its exact SU(3) value.

<sup>17</sup> This BATLEY 07 result uses our 2006 value of  $V_{ud}$  from semileptonic kaon decays as input.

## $g_2(0)/f_1(0)$ FOR $\Xi^0 \rightarrow \Sigma^+ e^- \bar{\nu}_e$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-1.7^{+2.1}_{-2.0} \pm 0.5</math></b>	487	<sup>18</sup> ALAVI-HARATI01i	KTEV	$p$ nucleus, 800 GeV

<sup>18</sup> ALAVI-HARATI 01i thus assumes that  $g_2 = 0$  in calculating  $g_2/f_1$ , above.

## $g_2(0)/f_1(0)$ FOR $\Xi^0 \rightarrow \Sigma^+ e^- \bar{\nu}_e$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.0 \pm 0.9</math> OUR AVERAGE</b>				
$2.0 \pm 1.3$		BATLEY	13 NA48	$p$ Be, 400 GeV
$2.0 \pm 1.2 \pm 0.5$	487	ALAVI-HARATI01i	KTEV	$p$ nucleus, 800 GeV

## $\Xi^0$ REFERENCES

BATLEY	13	PL B720 105	J.R. Batley <i>et al.</i>	(CERN NA48/1 Collab.)
BATLEY	10B	PL B693 241	J.R. Batley <i>et al.</i>	(CERN NA48/1 Collab.)
BATLEY	07	PL B645 36	J.R. Batley <i>et al.</i>	(CERN NA48/1 Collab.)
BATLEY	07C	PL B650 1	J.R. Batley <i>et al.</i>	(CERN NA48 Collab.)
ABOUZAID	05	PRL 95 081801	E. Abouzaid <i>et al.</i>	(FNAL KTeV Collab.)
WHITE	05	PRL 94 101804	C.G. White <i>et al.</i>	(FNAL HyperCP Collab.)
LAI	04A	PL B584 251	A. Lai <i>et al.</i>	(CERN NA48 Collab.)
JAMES	02	PRL 89 169901 (errat.)	C. James <i>et al.</i>	(MINN, MICH, WISC, RUTG)
TEIGE	02	PRL 89 169902 (errat.)	S. Teige <i>et al.</i>	(RUTG, MICH, MINN)
ALAVI-HARATI 01C		PRL 86 3239	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
ALAVI-HARATI 01i		PRL 87 132001	A. Alavi-Harati <i>et al.</i>	(FNAL KTeV Collab.)
FANTI	00	EPJ C12 69	V. Fanti <i>et al.</i>	(CERN NA48 Collab.)
AFFOLDER	99	PRL 82 3751	A. Affolder <i>et al.</i>	(FNAL KTeV Collab.)
JAMES	90	PRL 64 843	C. James <i>et al.</i>	(MINN, MICH, WISC, RUTG)
TEIGE	89	PRL 63 2717	S. Teige <i>et al.</i>	(RUTG, MICH, MINN)
HANDLER	82	PR D25 639	R. Handler <i>et al.</i>	(WISC, MICH, MINN+)
COX	81	PRL 46 877	P.T. Cox <i>et al.</i>	(MICH, WISC, RUTG, MINN+)
BUNCE	79	PL B6B 386	G.R.M. Bunce <i>et al.</i>	(G.R.M. Bunce <i>et al.</i> )
BUNCE	78	PR D18 636	G.R.M. Bunce <i>et al.</i>	(WISC, MICH, RUTG)
ZECH	77	NP B124 413	G. Zech <i>et al.</i>	(SIEG, CERN, DORT, HEIDH)
GEWENIGER	75	PL 57B 193	C. Geweniger <i>et al.</i>	(CERN, HEIDH)
BALTAY	74	PR D9 49	C. Baltay <i>et al.</i>	(COLU, BING, J)
YEH	74	PR D10 3545	N. Yeh <i>et al.</i>	(BING, COLU)
MAYEUR	72	NP B47 333	C. Mayeur <i>et al.</i>	(BRUX, CERN, TUFTS, LOUC)
Also		NP B53 268 (erratum)	C. Mayeur	
WILQUET	72	PL 42B 372	G. Wilquet <i>et al.</i>	(BRUX, CERN, TUFTS+)
DAUBER	69	PR 179 1262	P.M. Dauber <i>et al.</i>	(LRL)
PALMER	68	PL 26B 323	R.B. Palmer <i>et al.</i>	(BNL, SYRA)
BERGE	66	PR 147 945	J.P. Berge <i>et al.</i>	(LRL)
HUBBARD	66	Thesis UCRL 11510	J.R. Hubbard	(LRL)
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA)
PJERROU	65B	PRL 14 275	G.M. Pjerrou <i>et al.</i>	(UCLA)
Also		Thesis	G.M. Pjerrou	(UCLA)
CARMONY	64B	PRL 12 482	D.D. Carmony <i>et al.</i>	(UCLA)
HUBBARD	64	PR 135 B183	J.R. Hubbard <i>et al.</i>	(LRL)
JAUNEAU	63	PL 4 49	L. Jauneau <i>et al.</i>	(EPOL, CERN, LOUC+)
Also		Siena Conf. 1 1	L. Jauneau <i>et al.</i>	(EPOL, CERN, LOUC+)



$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+) \text{ Status: } ***$$

The parity has not actually been measured, but + is of course expected.

We have omitted some results that have been superseded by later experiments. See our earlier editions.

## $\Xi^-$ MASS

The fit uses the  $\Xi^-$ ,  $\Xi^+$ , and  $\Xi^0$  masses and the  $\Xi^- - \Xi^+$  mass difference. It assumes that the  $\Xi^-$  and  $\Xi^+$  masses are the same.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1321.71 \pm 0.07</math> OUR FIT</b>				
<b><math>1321.70 \pm 0.08 \pm 0.05</math></b>	2478 $\pm$ 68	ABDALLAH	06E DLPH	from Z decays
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1321.46 \pm 0.34$	632	DIBIANCA	75 DBC	4.9 GeV/c $K^- p$
$1321.12 \pm 0.41$	268	WILQUET	72 HLBC	
$1321.87 \pm 0.51$	195	<sup>1</sup> GOLDWASSER 70	HBC	5.5 GeV/c $K^- p$
$1321.67 \pm 0.52$	6	CHIEN	66 HBC	6.9 GeV/c $\bar{p} p$
$1321.4 \pm 1.1$	299	LONDON	66 HBC	
$1321.3 \pm 0.4$	149	PJERROU	65B HBC	
$1321.1 \pm 0.3$	241	<sup>2</sup> BADIER	64 HBC	
$1321.4 \pm 0.4$	517	<sup>2</sup> JAUNEAU	63D FBC	
$1321.1 \pm 0.65$	62	<sup>2</sup> SCHNEIDER	63 HBC	

<sup>1</sup> GOLDWASSER 70 uses  $m_{\Lambda} = 1115.58$  MeV.

<sup>2</sup> These masses have been increased 0.09 MeV because the  $\Lambda$  mass increased.

## $\Xi^+$ MASS

The fit uses the  $\Xi^-$ ,  $\Xi^+$ , and  $\Xi^0$  masses and the  $\Xi^- - \Xi^+$  mass difference. It assumes that the  $\Xi^-$  and  $\Xi^+$  masses are the same.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1321.71 \pm 0.07</math> OUR FIT</b>				
<b><math>1321.73 \pm 0.08 \pm 0.05</math></b>	2256 $\pm$ 63	ABDALLAH	06E DLPH	from Z decays
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$1321.6 \pm 0.8$	35	VOTRUBA	72 HBC	10 GeV/c $K^+ p$
$1321.2 \pm 0.4$	34	STONE	70 HBC	
$1320.69 \pm 0.93$	5	CHIEN	66 HBC	6.9 GeV/c $\bar{p} p$

$$(m_{\Xi^-} - m_{\Xi^+}) / m_{\Xi^-}$$

A test of  $CPT$  invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>(-2.5 \pm 8.7) \times 10^{-5}</math></b>	ABDALLAH	06E DLPH	from Z decays

Ξ<sup>-</sup> MEAN LIFE

Measurements with an error > 0.2 × 10<sup>-10</sup> s or with systematic errors not included have been omitted.

VALUE (10 <sup>-10</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.639±0.015 OUR AVERAGE</b>				
1.65 ±0.07 ±0.12	2478 ±68	ABDALLAH	06E	DLPH from Z decays
1.652±0.051	32k	BOURQUIN	84	SPEC Hyperon beam
1.665±0.065	41k	BOURQUIN	79	SPEC Hyperon beam
1.609±0.028	4286	HEMINGWAY	78	HBC 4.2 GeV/c K <sup>-</sup> p
1.67 ±0.08		DIBIANCA	75	DBC 4.9 GeV/c K <sup>-</sup> d
1.63 ±0.03	4303	BALTAY	74	HBC 1.75 GeV/c K <sup>-</sup> p
1.73 <sup>+0.08</sup> <sub>-0.07</sub>	680	MAYEUR	72	HLBC 2.1 GeV/c K <sup>-</sup>
1.61 ±0.04	2610	DAUBER	69	HBC
1.80 ±0.16	299	LONDON	66	HBC
1.70 ±0.12	246	PJERROU	65B	HBC
1.69 ±0.07	794	HUBBARD	64	HBC
1.86 <sup>+0.15</sup> <sub>-0.14</sub>	517	JAUNEAU	63D	FBC

Ξ<sup>+</sup> MEAN LIFE

VALUE (10 <sup>-10</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.70±0.08±0.12</b>	2256 ±63	ABDALLAH	06E	DLPH from Z decays
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.55 <sup>+0.35</sup> <sub>-0.20</sub>	35	<sup>3</sup> VOTRUBA	72	HBC 10 GeV/c K <sup>+</sup> p
1.6 ±0.3	34	STONE	70	HBC
1.9 <sup>+0.7</sup> <sub>-0.5</sub>	12	<sup>3</sup> SHEN	67	HBC
1.51±0.55	5	<sup>3</sup> CHIEN	66	HBC 6.9 GeV/c $\overline{p}$ p

<sup>3</sup> The error is statistical only.

$$(\tau_{\Xi^-} - \tau_{\Xi^+}) / \tau_{\Xi^-}$$

A test of CPT invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.01±0.07</b>	ABDALLAH	06E	DLPH from Z decays

Ξ<sup>-</sup> MAGNETIC MOMENT

See the "Note on Baryon Magnetic Moments" in the Λ Listings.

VALUE (μ <sub>N</sub> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.6507±0.0025 OUR AVERAGE</b>				
-0.6505±0.0025	4.36M	DURYEA	92	SPEC 800 GeV p Be
-0.661 ±0.036 ±0.036	44k	TROST	89	SPEC Ξ <sup>-</sup> ~ 250 GeV
-0.69 ±0.04	218k	RAMEIKA	84	SPEC 400 GeV pBe
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.674 ±0.021 ±0.020	122k	HO	90	SPEC See DURYEA 92
-2.1 ±0.8	2436	COOL	74	OSPK 1.8 GeV/c K <sup>-</sup> p
-0.1 ±2.1	2724	BINGHAM	70B	OSPK 1.8 GeV/c K <sup>-</sup> p

Ξ<sup>+</sup> MAGNETIC MOMENT

See the "Note on Baryon Magnetic Moments" in the Λ Listings.

VALUE (μ <sub>N</sub> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.657±0.028±0.020</b>	70k	HO	90	SPEC 800 GeV pBe

$$(\mu_{\Xi^-} + \mu_{\Xi^+}) / |\mu_{\Xi^-}|$$

A test of CPT invariance. We calculate this from the Ξ<sup>-</sup> and Ξ<sup>+</sup> magnetic moments above.

VALUE	DOCUMENT ID
<b>+0.01±0.05 OUR EVALUATION</b>	

Ξ<sup>-</sup> DECAY MODES

Mode	Fraction (Γ <sub>i</sub> /Γ)	Confidence level
Γ <sub>1</sub> Λπ <sup>-</sup>	(99.887±0.035) %	
Γ <sub>2</sub> Σ <sup>-</sup> γ	( 1.27 ±0.23 ) × 10 <sup>-4</sup>	
Γ <sub>3</sub> Λe <sup>-</sup> $\overline{\nu}_e$	( 5.63 ±0.31 ) × 10 <sup>-4</sup>	
Γ <sub>4</sub> Λμ <sup>-</sup> $\overline{\nu}_\mu$	( 3.5 <sup>+3.5</sup> <sub>-2.2</sub> ) × 10 <sup>-4</sup>	
Γ <sub>5</sub> Σ <sup>0</sup> e <sup>-</sup> $\overline{\nu}_e$	( 8.7 ±1.7 ) × 10 <sup>-5</sup>	
Γ <sub>6</sub> Σ <sup>0</sup> μ <sup>-</sup> $\overline{\nu}_\mu$	< 8	× 10 <sup>-4</sup> 90%
Γ <sub>7</sub> Ξ <sup>0</sup> e <sup>-</sup> $\overline{\nu}_e$	< 2.3	× 10 <sup>-3</sup> 90%

ΔS = 2 forbidden (S2) modes

Γ <sub>8</sub> nπ <sup>-</sup>	S2	< 1.9	× 10 <sup>-5</sup>	90%
Γ <sub>9</sub> n e <sup>-</sup> $\overline{\nu}_e$	S2	< 3.2	× 10 <sup>-3</sup>	90%
Γ <sub>10</sub> n μ <sup>-</sup> $\overline{\nu}_\mu$	S2	< 1.5	%	90%
Γ <sub>11</sub> pπ <sup>-</sup> π <sup>-</sup>	S2	< 4	× 10 <sup>-4</sup>	90%
Γ <sub>12</sub> pπ <sup>-</sup> e <sup>-</sup> $\overline{\nu}_e$	S2	< 4	× 10 <sup>-4</sup>	90%
Γ <sub>13</sub> pπ <sup>-</sup> μ <sup>-</sup> $\overline{\nu}_\mu$	S2	< 4	× 10 <sup>-4</sup>	90%
Γ <sub>14</sub> p μ <sup>-</sup> μ <sup>-</sup>	L	< 4	× 10 <sup>-8</sup>	90%

CONSTRAINED FIT INFORMATION

An overall fit to 4 branching ratios uses 5 measurements and one constraint to determine 5 parameters. The overall fit has a χ<sup>2</sup> = 1.0 for 1 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients <δx<sub>i</sub>δx<sub>j</sub>>/(δx<sub>i</sub>δx<sub>j</sub>), in percent, from the fit to the branching fractions, x<sub>i</sub> ≡ Γ<sub>i</sub>/Γ<sub>total</sub>. The fit constrains the x<sub>i</sub> whose labels appear in this array to sum to one.

x <sub>2</sub>	-6			
x <sub>3</sub>	-8	0		
x <sub>4</sub>	-99	0	-1	
x <sub>5</sub>	-5	0	0	0
	x <sub>1</sub>	x <sub>2</sub>	x <sub>3</sub>	x <sub>4</sub>

Ξ<sup>-</sup> BRANCHING RATIOS

A number of early results have been omitted.

$\Gamma(\Sigma^-\gamma)/\Gamma(\Lambda\pi^-)$					$\Gamma_2/\Gamma_1$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.27±0.24 OUR FIT</b>					
<b>1.27±0.23 OUR AVERAGE</b>					
1.22±0.23±0.06	211	<sup>4</sup> DUBBS	94	E761	$\Xi^-$ 375 GeV
2.27±1.02	9	BIAGI	87B	SPEC	SPS hyperon beam
4 DUBBS 94 also finds weak evidence that the asymmetry parameter $\alpha_\gamma$ is positive ( $\alpha_\gamma = 1.0 \pm 1.3$ ).					

<sup>4</sup> DUBBS 94 also finds weak evidence that the asymmetry parameter α<sub>γ</sub> is positive (α<sub>γ</sub> = 1.0 ± 1.3).

$\Gamma(\Lambda e^- \bar{\nu}_e)/\Gamma(\Lambda \pi^-)$				$\Gamma_3/\Gamma_1$
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.564±0.031 OUR FIT</b>				
<b>0.564±0.031</b>	2857	BOURQUIN	83	SPEC SPS hyperon beam
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.30 ±0.13	11	THOMPSON	80	ASPK Hyperon beam

$\Gamma(\Lambda\mu^-\overline{\nu}_\mu)/\Gamma(\Lambda\pi^-)$			$\Gamma_4/\Gamma_1$		
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$0.35^{+0.35}_{-0.22}$ OUR FIT					
$0.35\pm0.35$	1	YEH	74	HBC	Effective denom.=2859
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 2.3	90	0	THOMPSON	80	ASPK Effective denom.=1017
< 1.3			DAUBER	69	HBC
<12			BERGE	66	HBC

$\Gamma(\Sigma^0 e^- \bar{\nu}_e)/\Gamma(\Lambda \pi^-)$	$\Gamma_5/\Gamma_1$				
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.087±0.017 OUR FIT</b>					
<b>0.087±0.017</b>	154	BOURQUIN	83	SPEC	SPS hyperon beam

$[\Gamma(\Lambda e^- \overline{\nu}_e) + \Gamma(\Sigma^0 e^- \overline{\nu}_e)] / \Gamma(\Lambda \pi^-)$	$(\Gamma_3 + \Gamma_5) / \Gamma_1$			
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.651 \pm 0.031$	3011	<sup>5</sup> BOURQUIN	83	SPEC SPS hyperon beam
$0.68 \pm 0.22$	17	<sup>6</sup> DUCLOS	71	OSPK

<sup>5</sup> See the separate BOURQUIN 83 values for  $\Gamma(\Lambda e^- \overline{\nu}_e) / \Gamma(\Lambda \pi^-)$  and  $\Gamma(\Sigma^0 e^- \overline{\nu}_e) / \Gamma(\Lambda \pi^-)$  above.

<sup>6</sup> DUCLOS 71 cannot distinguish  $\Sigma^0$ 's from  $\Lambda$ 's. The Cabibbo theory predicts the  $\Sigma^0$  rate is about a factor 6 smaller than the  $\Lambda$  rate.

<sup>5</sup> See the separate BOURQUIN 83 values for Γ(Λe<sup>-</sup>  $\overline{\nu}_e$ )/Γ(Λπ<sup>-</sup>) and Γ(Σ<sup>0</sup> e<sup>-</sup>  $\overline{\nu}_e$ )/Γ(Λπ<sup>-</sup>) above.

<sup>6</sup> DUCLOS 71 cannot distinguish Σ<sup>0</sup>s from Λ's. The Cabibbo theory predicts the Σ<sup>0</sup> rate is about a factor 6 smaller than the Λ rate.

$\Gamma(\Sigma^0 \mu^- \bar{\nu}_\mu)/\Gamma(\Lambda \pi^-)$					$\Gamma_6/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.76</b>	90	0	YEH	74	HBC Effective denom.=3026
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<5			BERGE	66	HBC

$\Gamma(\Xi^0 e^- \bar{\nu}_e)/\Gamma(\Lambda \pi^-)$					$\Gamma_7/\Gamma_1$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<2.3	90	0	YEH	74	HBC Effective denom.=1000

# Baryon Particle Listings

Ξ<sup>-</sup>

$\Gamma(n\pi^-)/\Gamma(\Lambda\pi^-)$					$\Gamma_8/\Gamma_1$	
$\Delta S=2$ . Forbidden in first-order weak interaction.						
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.019</b>	90		BIAGI	82b	SPEC	SPS hyperon beam
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<3.0	90	0	YEH	74	HBC	Effective denom.=760
<1.1			DAUBER	69	HBC	
<5.0			FERRO-LUZZI	63	HBC	

$\Gamma(ne^- \bar{\nu}_e)/\Gamma(\Lambda\pi^-)$					$\Gamma_9/\Gamma_1$	
$\Delta S=2$ . Forbidden in first-order weak interaction.						
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 3.2</b>	90	0	YEH	74	HBC	Effective denom.=715
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<10	90		BINGHAM	65	RVUE	

$\Gamma(n\mu^-\bar{\nu}_\mu)/\Gamma(\Lambda\pi^-)$					$\Gamma_{10}/\Gamma_1$	
$\Delta S=2$ . Forbidden in first-order weak interaction.						
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt;15.3</b>	90	0	YEH	74	HBC	Effective denom.=150

$\Gamma(\rho\pi^-\pi^-)/\Gamma(\Lambda\pi^-)$					$\Gamma_{11}/\Gamma_1$	
$\Delta S=2$ . Forbidden in first-order weak interaction.						
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.7</b>	90	0	YEH	74	HBC	Effective denom.=6200

$\Gamma(\rho\pi^-e^-\bar{\nu}_e)/\Gamma(\Lambda\pi^-)$					$\Gamma_{12}/\Gamma_1$	
$\Delta S=2$ . Forbidden in first-order weak interaction.						
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.7</b>	90	0	YEH	74	HBC	Effective denom.=6200

$\Gamma(\rho\pi^-\mu^-\bar{\nu}_\mu)/\Gamma(\Lambda\pi^-)$					$\Gamma_{13}/\Gamma_1$	
$\Delta S=2$ . Forbidden in first-order weak interaction.						
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.7</b>	90	0	YEH	74	HBC	Effective denom.=6200

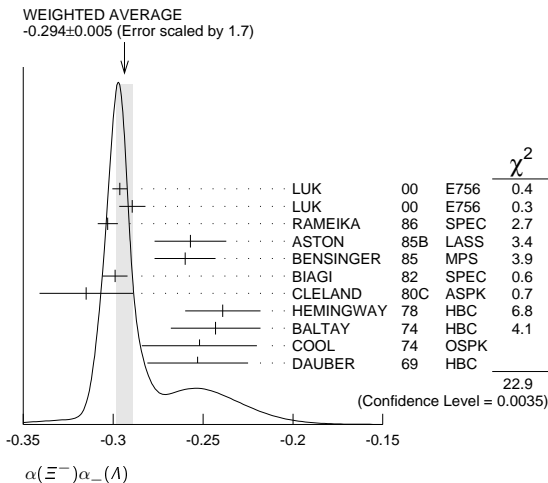
$\Gamma(\rho\mu^-\mu^-)/\Gamma(\Lambda\pi^-)$					$\Gamma_{14}/\Gamma_1$	
$\Delta L=2$ decay, forbidden by total lepton number conservation.						
VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	COMMENT		
<b>&lt;4.0</b>	90	RAJARAM	05	HYCP	$p$ Cu, 800 GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$<3.7 \times 10^4$	90	<sup>7</sup> LITTENBERG	92b	HBC	Uses YEH 74 data	
<sup>7</sup> This LITTENBERG 92b limit and the identical YEH 74 limits for the preceding three modes all result from nonobservance of any 3-prong decays of the $\Xi^-$ . One could as well apply the limit to the <i>sum</i> of the four modes.						

## Ξ<sup>-</sup> DECAY PARAMETERS

See the “Note on Baryon Decay Parameters” in the neutron Listings.

$\alpha(\Xi^-)\alpha_-(\Lambda)$				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.294 ± 0.005 OUR AVERAGE</b>				Error includes scale factor of 1.7. See the ideogram below.
-0.2963 ± 0.0042	189k	LUK	00	E756 $p$ Be, 800 GeV
-0.2894 ± 0.0073	63k	LUK	00	E756 $p$ Be, 800 GeV
-0.303 ± 0.004 ± 0.004	192k	RAMEIKA	86	SPEC 400 GeV $p$ Be
-0.257 ± 0.020	11k	ASTON	85b	LASS 11 GeV/c $K^-\pi$
-0.260 ± 0.017	21k	BENSINGER	85	MPS 5 GeV/c $K^-\pi$
-0.299 ± 0.007	150k	BIAGI	82	SPEC SPS hyperon beam
-0.315 ± 0.026	9046	CLELAND	80c	ASPK BNL hyperon beam
-0.239 ± 0.021	6599	HEMINGWAY	78	HBC 4.2 GeV/c $K^-\pi$
-0.243 ± 0.025	4303	BALTAY	74	HBC 1.75 GeV/c $K^-\pi$
-0.252 ± 0.032	2436	COOL	74	OSPK 1.8 GeV/c $K^-\pi$
-0.253 ± 0.028	2781	DAUBER	69	HBC

8 This LUK 00 value is for  $\alpha(\Xi^+)\alpha_+(\bar{\Lambda})$ . We assume  $CP$  conservation here by including it in the average for  $\alpha(\Xi^-)\alpha_-(\Lambda)$ . But see the second data block below for the  $CP$  test.



**α FOR Ξ<sup>-</sup> → Λπ<sup>-</sup>**  
The above average,  $\alpha(\Xi^-)\alpha_-(\Lambda) = -0.294 \pm 0.005$ , where the error includes a scale factor of 1.7, divided by our current average  $\alpha_-(\Lambda) = 0.642 \pm 0.013$ , gives the following value for  $\alpha(\Xi^-)$ .

VALUE	DOCUMENT ID
<b>-0.458 ± 0.012 OUR EVALUATION</b>	Error includes scale factor of 1.8.

**$\frac{[\alpha(\Xi^-)\alpha_-(\Lambda) - \alpha(\Xi^+)\alpha_+(\bar{\Lambda})]}{[\alpha(\Xi^-)\alpha_-(\Lambda) + \alpha(\Xi^+)\alpha_+(\bar{\Lambda})]}$**   
This is zero if  $CP$  is conserved. The  $\alpha$ 's are the decay-asymmetry parameters for  $\Xi^- \rightarrow \Lambda\pi^-$  and  $\Lambda \rightarrow p\pi^-$  and for  $\Xi^+ \rightarrow \bar{\Lambda}\pi^+$  and  $\bar{\Lambda} \rightarrow \bar{p}\pi^+$ .

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0 ± 5.1 ± 4.4</b>	158M	HOLMSTROM 04	HYCP	$p$ Cu, 800 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
+120 ± 140	252k	LUK	00	E756 $p$ Be, 800 GeV

**φ ANGLE FOR Ξ<sup>-</sup> → Λπ<sup>-</sup>** ( $\tan\phi = \beta/\gamma$ )

VALUE (°)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>- 2.1 ± 0.8 OUR AVERAGE</b>				
- 2.39 ± 0.64 ± 0.64	144M	<sup>9</sup> HUANG	04	HYCP $p$ Cu, 800 GeV
- 1.61 ± 2.66 ± 0.37	1.35M	<sup>10</sup> CHAKRAVO...	03	E756 $p$ Be, 800 GeV
5 ± 10	11k	ASTON	85b	LASS $K^-\pi$
14.7 ± 16.0	21k	<sup>11</sup> BENSINGER	85	MPS 5 GeV/c $K^-\pi$
11 ± 9	4303	BALTAY	74	HBC 1.75 GeV/c $K^-\pi$
5 ± 16	2436	COOL	74	OSPK 1.8 GeV/c $K^-\pi$
-14 ± 11	2781	DAUBER	69	HBC Uses $\alpha_\Lambda = 0.647 \pm 0.020$
0 ± 12	1004	<sup>12</sup> BERGE	66	HBC
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-26 ± 30	2724	BINGHAM	70b	OSPK
0 ± 20.4	364	<sup>12</sup> LONDON	66	HBC Using $\alpha_\Lambda = 0.62$
54 ± 30	356	<sup>12</sup> CARMONY	64b	HBC

<sup>9</sup> From this result and  $\alpha_\Xi$ , HUANG 04 gets  $\beta_\Xi = -0.037 \pm 0.011 \pm 0.010$  and  $\gamma_\Xi = 0.888 \pm 0.0004 \pm 0.006$ . And the strong p-s phase difference for  $\Lambda\pi^-$  scattering is  $(4.6 \pm 1.4 \pm 1.2)^\circ$ .

<sup>10</sup> From this result and  $\alpha_\Xi$ , CHAKRAVORTY 03 obtains  $\beta_\Xi = -0.025 \pm 0.042 \pm 0.006$  and  $\gamma_\Xi = 0.889 \pm 0.001 \pm 0.007$ . And the strong p-s phase difference for  $\Lambda\pi^-$  scattering is  $(3.17 \pm 5.28 \pm 0.73)^\circ$ .

<sup>11</sup> BENSINGER 85 used  $\alpha_\Lambda = 0.642 \pm 0.013$ .

<sup>12</sup> The errors have been multiplied by 1.2 due to approximations used for the  $\Xi$  polarization; see DAUBER 69 for a discussion.

**g<sub>A</sub> / g<sub>V</sub> FOR Ξ<sup>-</sup> → Λe<sup>-</sup>ν<sub>e</sub>**

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.25 ± 0.05</b>	1992	<sup>13</sup> BOURQUIN	83	SPEC SPS hyperon beam

<sup>13</sup> BOURQUIN 83 assumes that  $g_2 = 0$ . Also, the sign has been changed to agree with our conventions, given in the “Note on Baryon Decay Parameters” in the neutron Listings.

## Ξ<sup>-</sup> REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

ABDALLAH	06E	PL B639 179	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
RAJARAM	05	PRL 94 181801	D. Rajaram <i>et al.</i>	(FNAL HyperCP Collab.)
HOLMSTROM	04	PRL 93 262001	T. Holmstrom <i>et al.</i>	(FNAL HyperCP Collab.)
HUANG	04	PRL 93 011802	M. Huang <i>et al.</i>	(FNAL HyperCP Collab.)
CHAKRAVO...	03	PRL 91 031601	A. Chakravorty <i>et al.</i>	(FNAL E756 Collab.)
LUK	00	PRL 85 4860	K.B. Luk <i>et al.</i>	(FNAL E756 Collab.)
DUBBS	94	PRL 72 8008	T. Dubbs <i>et al.</i>	(FNAL E761 Collab.)
DURVEA	92	PRL 68 7668	J. Durvea <i>et al.</i>	(MINN, FNAL, MICH, RUTG)
LITTENBERG	92b	PR D46 R892	L.S. Littenberg, R.E. Shrock	(BNL, STON)
HO	90	PRL 65 1713	P.M. Ho <i>et al.</i>	(MICH, FNAL, MINN, RUTG)
Also		PR D44 3402	P.M. Ho <i>et al.</i>	(MICH, FNAL, MINN, RUTG)
TROST	89	PR D40 1703	L.H. Trost <i>et al.</i>	(FNAL-715 Collab.)
BIAGI	87b	ZPHY C35 143	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
RAMEIKA	86	PR D33 3172	R. Rameika <i>et al.</i>	(RUTG, MICH, WISC+)
ASTON	85b	PR D32 2270	D. Aston <i>et al.</i>	(SLAC, CARL, CNRC, CINC)
BENSINGER	85	NP B252 561	J.R. Bensinger <i>et al.</i>	(CHIC, ELMT, FNAL+)
BOURQUIN	84	NP B241 1	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
RAMEIKA	84	PRL 52 581	R. Rameika <i>et al.</i>	(RUTG, MICH, WISC+)
BOURQUIN	83	ZPHY C21 1	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
BIAGI	82	PL 112B 265	S.F. Biagi <i>et al.</i>	(BRIS, CAVE, GEVA+)
BIAGI	82b	PL 112B 277	S.F. Biagi <i>et al.</i>	(LOQM, GEVA, RL+)
CLELAND	80c	PR D21 12	W.E. Cleland <i>et al.</i>	(PITT, BNL)
THOMPSON	80	PR D21 25	J.A. Thompson <i>et al.</i>	(PITT, BNL)
BOURQUIN	79	PL 87B 297	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
HEMINGWAY	78	NP B142 205	R.J. Hemingway <i>et al.</i>	(CERN, ZEEM, NJM+)
DIBIANCA	75	NP B98 137	F.A. Dibiaccia, R.J. Endorf	(CMU)
BALTAY	74	PR D9 49	C. Baltay <i>et al.</i>	(COLU, BING) J
COOL	74	PR D10 792	R.L. Cool <i>et al.</i>	(BNL)
Also		PR D10 3545	R.L. Cool <i>et al.</i>	(BNL)
YEH	74	PR D10 3545	N. Yeh <i>et al.</i>	(BING, COLU)
MAYEUR	72	NP B47 333	C. Mayeur <i>et al.</i>	(BRUX, CERN, TUFTS, LOUC)
VOTRUBA	72	NP B45 77	M.F. Votruba, A. Saifer, T.M. Ratcliffe	(BIRM+)
WILQUET	72	PL 42B 372	G. Wilquet <i>et al.</i>	(BRUX, CERN, TUFTS+)
DUCLOS	71	NP B32 493	J. Duclos <i>et al.</i>	(CERN)
BINGHAM	70b	PR D1 3010	G.M. Bingham <i>et al.</i>	(UCSD, WASH)
GOLDWASSER	70	PR D1 1960	E.L. Goldwasser, P.F. Schultz	(ILL)
STONE	70	PL 32B 515	S.L. Stone <i>et al.</i>	(ROCH)
DAUBER	69	PR 179 1262	P.M. Dauber <i>et al.</i>	(LRL) J
SHEN	67	PL 25B 443	B.C. Shen, A. Firestone, G. Goldhaber	(UCB+)



BERGE	66	PR 147 945	J.P. Berge <i>et al.</i>	(LRL)
CHIEN	66	PR 152 1171	C.Y. Chien <i>et al.</i>	(YALE, BNL)
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA)
BINGHAM	65	PRSL 285 202	H.H. Bingham	(CERN)
PJERROU	65B	PRL 14 275	G.M. Pjerrou <i>et al.</i>	(UCLA)
Also		Thesis	G.M. Pjerrou	(UCLA)
BADIER	64	Dubna Conf. 1 593	J. Badier <i>et al.</i>	(EPOL, SACL, ZEEM)
CARMONY	64B	PRL 12 482	D.D. Carmony <i>et al.</i>	(UCLA) J
HUBBARD	64	PR 135 B183	J.R. Hubbard <i>et al.</i>	(LRL)
FERRO-LUZZI	63	PR 130 1568	M. Ferro-Luzzi <i>et al.</i>	(LRL)
JAUNEAU	63D	Siena Conf. 4	L. Jauneau <i>et al.</i>	(EPOL, CERN, LOUC+)
Also		PL 5 261	L. Jauneau <i>et al.</i>	(EPOL, CERN, LOUC+)
SCHNEIDER	63	PL 4 360	J. Schneider	(CERN)

## $\Xi$ RESONANCES

The accompanying table gives our evaluation of the present status of the  $\Xi$  resonances. Not much is known about  $\Xi$  resonances. This is because (1) they can only be produced as a part of a final state, and so the analysis is more complicated than if direct formation were possible, (2) the production cross sections are small (typically a few  $\mu\text{b}$ ), and (3) the final states are topologically complicated and difficult to study with electronic techniques. Thus early information about  $\Xi$  resonances came entirely from bubble chamber experiments, where the numbers of events are small, and only in the 1980's did electronic experiments make any significant contributions. However, nothing of significance on  $\Xi$  resonances has been added since our 1988 edition.

For a detailed earlier review, see Meadows [1].

Table 1. The status of the  $\Xi$  resonances. Only those with an overall status of \*\*\* or \*\*\*\* are included in the Baryon Summary Table.

Particle	$J^P$	Overall status	Status as seen in —				
			$\Xi\pi$	$\Lambda K$	$\Sigma K$	$\Xi(1530)\pi$	Other channels
$\Xi(1318)$	$1/2^+$	****					Decays weakly
$\Xi(1530)$	$3/2^+$	****	****				
$\Xi(1620)$		*	*				
$\Xi(1690)$		***		***	**		
$\Xi(1820)$	$3/2^-$	***	**	***	**	**	
$\Xi(1950)$		***	**	**		*	
$\Xi(2030)$		***		**	***		
$\Xi(2120)$		*		*			
$\Xi(2250)$		**					3-body decays
$\Xi(2370)$		**					3-body decays
$\Xi(2500)$		*		*	*		3-body decays
****	Existence is certain, and properties are at least fairly well explored.						
***	Existence ranges from very likely to certain, but further confirmation is desirable and/or quantum numbers, branching fractions, etc. are not well determined.						
**	Evidence of existence is only fair.						
*	Evidence of existence is poor.						

## Reference

1. B.T. Meadows, in *Proceedings of the IV<sup>th</sup> International Conference on Baryon Resonances* (Toronto, 1980), ed. N. Isgur, p. 283.

## $\Xi(1530) 3/2^+$

$$I(J^P) = \frac{1}{2}(\frac{3}{2}^+) \text{ Status: } ****$$

This is the only  $\Xi$  resonance whose properties are all reasonably well known. Assuming that the  $\Lambda_c^+$  has  $J^P = 1/2^+$ , AUBERT 08AK, in a study of  $\Lambda_c^+ \rightarrow \Xi^- \pi^+ K^+$ , finds conclusively that the spin of the  $\Xi(1530)^0$  is  $3/2$ . In conjunction with SCHLEIN 63B and BUTTON-SHAFFER 66, this proves also that the parity is +.

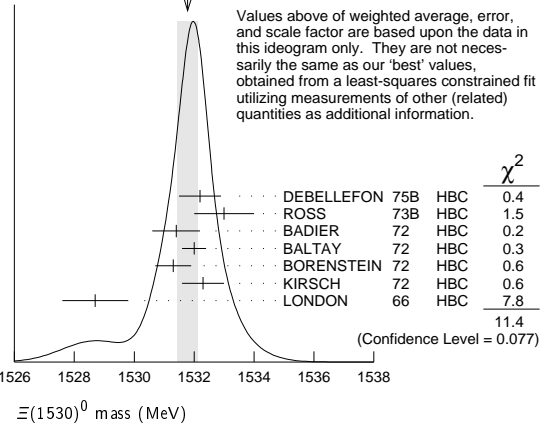
We use only those determinations of the mass and width that are accompanied by some discussion of systematics and resolution.

## $\Xi(1530)$ MASSES

### $\Xi(1530)^0$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1531.80 <math>\pm</math> 0.32 OUR FIT</b>				Error includes scale factor of 1.3.
<b>1531.78 <math>\pm</math> 0.34 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
1532.2 $\pm$ 0.7		DEBELLEFON 75B	HBC	$K^- p \rightarrow \Xi^- \bar{K} \pi$
1533 $\pm$ 1		ROSS 73B	HBC	$K^- p \rightarrow \Xi^- \bar{K} \pi(\pi)$
1531.4 $\pm$ 0.8	59	BADIER 72	HBC	$K^- p$ 3.95 GeV/c
1532.0 $\pm$ 0.4	1262	BALTAY 72	HBC	$K^- p$ 1.75 GeV/c
1531.3 $\pm$ 0.6	324	BORENSTEIN 72	HBC	$K^- p$ 2.2 GeV/c
1532.3 $\pm$ 0.7	286	KIRSCH 72	HBC	$K^- p$ 2.87 GeV/c
1528.7 $\pm$ 1.1	76	LONDON 66	HBC	$K^- p$ 2.24 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1532.1 $\pm$ 0.4	1244	ASTON 85B	LASS	$K^- p$ 11 GeV/c
1532.1 $\pm$ 0.6	2700	<sup>1</sup> BAUBILLIER 81B	HBC	$K^- p$ 8.25 GeV/c
1530 $\pm$ 1	450	BIAGI 81	SPEC	SPS hyperon beam
1527 $\pm$ 6	80	SIXEL 79	HBC	$K^- p$ 10 GeV/c
1535 $\pm$ 4	100	SIXEL 79	HBC	$K^- p$ 16 GeV/c
1533.6 $\pm$ 1.4	97	BERTHON 74	HBC	Quasi-2-body $\sigma$

WEIGHTED AVERAGE  
1531.78 $\pm$ 0.34 (Error scaled by 1.4)



### $\Xi(1530)^-$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1535.0 <math>\pm</math> 0.6 OUR FIT</b>				
<b>1535.2 <math>\pm</math> 0.8 OUR AVERAGE</b>				
1534.5 $\pm$ 1.2		DEBELLEFON 75B	HBC	$K^- p \rightarrow \Xi^- \bar{K} \pi$
1535.3 $\pm$ 2.0		ROSS 73B	HBC	$K^- p \rightarrow \Xi^- \bar{K} \pi(\pi)$
1536.2 $\pm$ 1.6	185	KIRSCH 72	HBC	$K^- p$ 2.87 GeV/c
1535.7 $\pm$ 3.2	38	LONDON 66	HBC	$K^- p$ 2.24 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1540 $\pm$ 3	48	BERTHON 74	HBC	Quasi-2-body $\sigma$
1534.7 $\pm$ 1.1	334	BALTAY 72	HBC	$K^- p$ 1.75 GeV/c

### $m_{\Xi(1530)^-} - m_{\Xi(1530)}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>3.2 <math>\pm</math> 0.6 OUR FIT</b>			
<b>2.9 <math>\pm</math> 0.9 OUR AVERAGE</b>			
2.7 $\pm$ 1.0	BALTAY 72	HBC	$K^- p$ 1.75 GeV/c
2.0 $\pm$ 3.2	MERRILL 66	HBC	$K^- p$ 1.7-2.7 GeV/c
5.7 $\pm$ 3.0	PJERROU 65B	HBC	$K^- p$ 1.8-1.95 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •			
3.9 $\pm$ 1.8	<sup>2</sup> KIRSCH 72	HBC	$K^- p$ 2.87 GeV/c
7 $\pm$ 4	<sup>2</sup> LONDON 66	HBC	$K^- p$ 2.24 GeV/c

## $\Xi(1530)$ WIDTHS

### $\Xi(1530)^0$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>9.1 <math>\pm</math> 0.5 OUR AVERAGE</b>			
9.5 $\pm$ 1.2	DEBELLEFON 75B	HBC	$K^- p \rightarrow \Xi^- \bar{K} \pi$
9.1 $\pm$ 2.4	ROSS 73B	HBC	$K^- p \rightarrow \Xi^- \bar{K} \pi(\pi)$
11 $\pm$ 2	BADIER 72	HBC	$K^- p$ 3.95 GeV/c
9.0 $\pm$ 0.7	BALTAY 72	HBC	$K^- p$ 1.75 GeV/c
8.4 $\pm$ 1.4	BORENSTEIN 72	HBC	$\Xi^- \pi^+$
11.0 $\pm$ 1.8	KIRSCH 72	HBC	$\Xi^- \pi^+$
7 $\pm$ 7	BERGE 66	HBC	$K^- p$ 1.5-1.7 GeV/c
8.5 $\pm$ 3.5	LONDON 66	HBC	$K^- p$ 2.24 GeV/c
7 $\pm$ 2	SCHLEIN 63B	HBC	$K^- p$ 1.8, 1.95 GeV/c

Baryon Particle Listings

Ξ(1530), Ξ(1620), Ξ(1690)

• • • We do not use the following data for averages, fits, limits, etc. • • •

12.8±1.0	2700	<sup>1</sup> BAUBILLIER	81B	HBC	$K^-p$ 8.25 GeV/c
19 ±6	80	<sup>3</sup> SIXEL	79	HBC	$K^-p$ 10 GeV/c
14 ±5	100	<sup>3</sup> SIXEL	79	HBC	$K^-p$ 16 GeV/c

Ξ(1530)<sup>-</sup> WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>9.9<sup>+1.7</sup><sub>-1.9</sub> OUR AVERAGE</b>			
9.6±2.8	DEBELLEFON	75B	HBC $K^-p \rightarrow \Xi^- \overline{K} \pi$
8.3±3.6	ROSS	73B	HBC $K^-p \rightarrow \Xi^- \overline{K} \pi (\pi)$
7.8 <sup>+3.5</sup> <sub>-7.8</sub>	BALTAY	72	HBC $K^-p$ 1.75 GeV/c
16.2±4.6	KIRSCH	72	HBC $\Xi^- \pi^0, \Xi^0 \pi^-$

Ξ(1530) POLE POSITIONS

Ξ(1530)<sup>0</sup> REAL PART

VALUE	DOCUMENT ID	COMMENT
1531.6±0.4	LICHTENBERG74	Using HABIBI 73

Ξ(1530)<sup>0</sup> IMAGINARY PART

VALUE	DOCUMENT ID	COMMENT
4.45±0.35	LICHTENBERG74	Using HABIBI 73

Ξ(1530)<sup>-</sup> REAL PART

VALUE	DOCUMENT ID	COMMENT
1534.4±1.1	LICHTENBERG74	Using HABIBI 73

Ξ(1530)<sup>-</sup> IMAGINARY PART

VALUE	DOCUMENT ID	COMMENT
3.9 <sup>+1.75</sup> <sub>-3.9</sub>	LICHTENBERG74	Using HABIBI 73

Ξ(1530) DECAY MODES

Mode	Fraction (Γ <sub>i</sub> /Γ)	Confidence level
Γ <sub>1</sub> Ξ π	100 %	
Γ <sub>2</sub> Ξ γ	<4 %	90%

Ξ(1530) BRANCHING RATIOS

Γ (Ξ γ)/Γ <sub>total</sub>	CL%	DOCUMENT ID	TECN	COMMENT	Γ <sub>2</sub> /Γ
<b>&lt;0.04</b>	90	KALBFLEISCH 75	HBC	$K^-p$ 2.18 GeV/c	

Ξ(1530) FOOTNOTES

- <sup>1</sup>BAUBILLIER 81B is a fit to the inclusive spectrum. The resolution (5 MeV) is not unfolded.  
<sup>2</sup>Redundant with data in the mass Listings.  
<sup>3</sup>SIXEL 79 doesn't unfold the experimental resolution of 15 MeV.

Ξ(1530) REFERENCES

AUBERT	08AK	PR D78 034008	B. Aubert <i>et al.</i>	(BABAR Collab.)
ASTON	85B	PR D32 2270	D. Aston <i>et al.</i>	(SLAC, CARL, CNRC, CINC)
BAUBILLIER	81B	NP B192 1	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)
BIAGI	81	ZPHY C9 305	S.F. Biagi <i>et al.</i>	(BRIS, CAVE, GEVA+)
SIXEL	79	NP B159 125	P. Sixel <i>et al.</i>	(AACH3, BERL, CERN, LOIC+)
DEBELLEFON	75B	NC 28A 289	A. de Bellefon <i>et al.</i>	(CDEF, SACL)
KALBFLEISCH	75	PR D11 987	G.R. Kalbfleisch, R.C. Strand, J.W. Chapman	(BNL+)
BERTHON	74	NC 21A 146	A. Berthon <i>et al.</i>	(CDEF, RHEL, SACL+)
LICHTENBERG	74	PR D10 3865	D.B. Lichtenberg	(IND)
Also		Private Comm.	D.B. Lichtenberg	(IND)
HABIBI	73	Thesis Nevis 199	M. Habibi	(COLU)
ROSS	73B	Purdue Conf. 355	R.T. Ross, J.L. Lloyd, D. Radojicic	(OXF)
BADIER	72	NP B37 429	J. Badier <i>et al.</i>	(EPOL)
BALTAY	72	PL 42B 129	C. Baltay <i>et al.</i>	(COLU, BING)
BORENSTEIN	72	PR D5 1559	S.R. Borenstein <i>et al.</i>	(BNL, MICH)I
KIRSCH	72	NP B40 349	L.E. Kirsch <i>et al.</i>	(BRAN, UMD, SYRA+)
BERGE	66	PR 147 945	J.P. Berge <i>et al.</i>	(LRL)I
BUTTON-...	66	PR 142 883	J. Button-Shafer <i>et al.</i>	(LRL)JP
LONDON	66	PR 143 1034	G.W. London <i>et al.</i>	(BNL, SYRA)IJ
MERRILL	66	Thesis UCRL 16455	D.W. Merrill	(LRL)JP
PJERROU	65B	PRL 14 275	G.M. Pjerrou <i>et al.</i>	(UCLA)
SCHLEIN	63B	PRL 11 167	P.E. Schlein <i>et al.</i>	(UCLA)IJP

OTHER RELATED PAPERS

MAZZUCATO	81	NP B178 1	M. Mazzucato <i>et al.</i>	(AMST, CERN, NIJM+)
BRIEFEL	77	PR D16 2706	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
BRIEFEL	75	PR D12 1859	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
HUNGERBU...	74	PR D10 2051	V. Hungerbuhler <i>et al.</i>	(YALE, FNAL, BNL+)
BUTTON-...	66	PR 142 883	J. Button-Shafer <i>et al.</i>	(LRL)JP

Ξ(1620)

$I(J^P) = \frac{1}{2}(?)^?$  Status: \*

*I, P* need confirmation.

OMITTED FROM SUMMARY TABLE

What little evidence there is consists of weak signals in the Ξπ channel. A number of other experiments (e.g., BORENSTEIN 72 and HASSALL 81) have looked for but not seen any effect.

Ξ(1620) MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>≈1620 OUR ESTIMATE</b>				
1624±3	31	BRIEFEL	77	HBC $K^-p$ 2.87 GeV/c
1633±12	34	DEBELLEFON	75B	HBC $K^-p \rightarrow \Xi^- \overline{K} \pi$
1606±6	29	ROSS	72	HBC $K^-p$ 3.1–3.7 GeV/c

Ξ(1620) WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
22.5	31	<sup>1</sup> BRIEFEL	77	HBC $K^-p$ 2.87 GeV/c
40 ±15	34	DEBELLEFON	75B	HBC $K^-p \rightarrow \Xi^- \overline{K} \pi$
21 ±7	29	ROSS	72	HBC $K^-p \rightarrow \Xi^- \pi^+ K^*0(892)$

Ξ(1620) DECAY MODES

Mode
Γ <sub>1</sub> Ξ π

Ξ(1620) FOOTNOTES

- <sup>1</sup>The fit is insensitive to values between 15 and 30 MeV.

Ξ(1620) REFERENCES

HASSALL	81	NP B189 397	J.K. Hassall <i>et al.</i>	(CAVE, MSU)
BRIEFEL	77	PR D16 2706	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
Also		Duke Conf. 317	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
Hyperon Resonances, 1970				
Also		PR D12 1859	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
DEBELLEFON	75B	NC 28A 289	A. de Bellefon <i>et al.</i>	(CDEF, SACL)
BORENSTEIN	72	PR D5 1559	S.R. Borenstein <i>et al.</i>	(BNL, MICH)I
ROSS	72	PL 38B 177	R.T. Ross <i>et al.</i>	(OXF)I

OTHER RELATED PAPERS

HUNGERBU...	74	PR D10 2051	V. Hungerbuhler <i>et al.</i>	(YALE, FNAL, BNL+)
SCHMIDT	73	Purdue Conf. 363	P.E. Schmidt	(BRAN)
KALBFLEISCH	70	Duke Conf. 331	G.R. Kalbfleisch	(BNL)I
Hyperon Resonances 1970				
APSELL	69	PRL 23 884	S.P. Apsell <i>et al.</i>	(BRAN, UMD, SYRA+)
BARTSCH	69	PL 28B 439	J. Bartsch <i>et al.</i>	(AACH, BERL, CERN+)

Ξ(1690)

$I(J^P) = \frac{1}{2}(?)^?$  Status: \*\*\*

AUBERT 08AK, in a study of  $\Lambda_c^+ \rightarrow \Xi^- \pi^+ K^+$ , finds some evidence that the Ξ(1690) has  $J^P = 1/2^-$ .

DIONISI 78 sees a threshold enhancement in both the neutral and negatively charged  $\Sigma^- \overline{K}$  mass spectra in  $K^-p \rightarrow (\Sigma^- \overline{K}) K \pi$  at 4.2 GeV/c. The data from the  $\Sigma^- \overline{K}$  channels alone cannot distinguish between a resonance and a large scattering length. Weaker evidence at the same mass is seen in the corresponding  $\Lambda^- \overline{K}$  channels, and a coupled-channel analysis yields results consistent with a new Ξ.

BIAGI 81 sees an enhancement at 1700 MeV in the diffractively produced  $\Lambda K^-$  system. A peak is also observed in the  $\Lambda \overline{K}^0$  mass spectrum at 1660 MeV that is consistent with a 1720 MeV resonance decaying to  $\Sigma^- \overline{K}^0$ , with the γ from the  $\Sigma^-$  decay not detected.

BIAGI 87 provides further confirmation of this state in diffractive dissociation of  $\Xi^-$  into  $\Lambda K^-$ . The significance claimed is 6.7 standard deviations.

ADAMOVIICH 98 sees a peak of 1400 ± 300 events in the  $\Xi^- \pi^+$  spectrum produced by 345 GeV/c Σ<sup>-</sup>-nucleus interactions.

Ξ(1690) MASSES

MIXED CHARGES

**1690±10 OUR ESTIMATE** This is only an educated guess; the error given is larger than the error on the average of the published values.

See key on page 601

# Baryon Particle Listings

## $\Xi(1690), \Xi(1820)$

### $\Xi(1690)^0$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
1686 ± 4	1400	ADAMOVICH	98 WA89	$\Sigma^-$ nucleus, 345 GeV/c
1699 ± 5	175	<sup>1</sup> DIONISI	78 HBC	$K^- p$ 4.2 GeV/c
1684 ± 5	183	<sup>2</sup> DIONISI	78 HBC	$K^- p$ 4.2 GeV/c

### $\Xi(1690)^-$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
1691.1 ± 1.9 ± 2.0	104	BIAGI	87 SPEC	$\Xi^-$ Be 116 GeV
1700 ± 10	150	<sup>3</sup> BIAGI	81 SPEC	$\Xi^-$ H 100, 135 GeV
1694 ± 6	45	<sup>4</sup> DIONISI	78 HBC	$K^- p$ 4.2 GeV/c

### $\Xi(1690)$ WIDTHS

#### MIXED CHARGES

VALUE (MeV)	DOCUMENT ID
<30 OUR ESTIMATE	

### $\Xi(1690)^0$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
10 ± 6	1400	ADAMOVICH	98 WA89	$\Sigma^-$ nucleus, 345 GeV/c
44 ± 23	175	<sup>1</sup> DIONISI	78 HBC	$K^- p$ 4.2 GeV/c
20 ± 4	183	<sup>2</sup> DIONISI	78 HBC	$K^- p$ 4.2 GeV/c

### $\Xi(1690)^-$ WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< 8	90	104	BIAGI	87 SPEC	$\Xi^-$ Be 116 GeV
47 ± 14		150	<sup>3</sup> BIAGI	81 SPEC	$\Xi^-$ H 100, 135 GeV
26 ± 6		45	<sup>4</sup> DIONISI	78 HBC	$K^- p$ 4.2 GeV/c

### $\Xi(1690)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\Lambda \bar{K}$	seen
$\Gamma_2$ $\Sigma \bar{K}$	seen
$\Gamma_3$ $\Xi \pi$	seen
$\Gamma_4$ $\Xi^- \pi^+ \pi^0$	
$\Gamma_5$ $\Xi^- \pi^+ \pi^-$	possibly seen
$\Gamma_6$ $\Xi(1530) \pi$	

### $\Xi(1690)$ BRANCHING RATIOS

$\Gamma(\Lambda\bar{K})/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
seen	104	BIAGI	87 SPEC	—	$\Xi^-$ Be 116 GeV

$\Gamma(\Sigma K)/\Gamma(\Lambda K)$					$\Gamma_2/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
0.75 ± 0.39	75	ABE	02c BELL		$e^+e^- \approx \Upsilon(4S)$
2.7 ± 0.9		DIONISI	78 HBC	0	$K^-p$ 4.2 GeV/c
3.1 ± 1.4		DIONISI	78 HBC	—	$K^-p$ 4.2 GeV/c

$\Gamma(\Xi\pi)/\Gamma(\Sigma\bar{K})$	$\Gamma_3/\Gamma_2$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<0.09	DIONISI	78	HBC	0	$K^-p$ 4.2 GeV/c

$\Gamma(\Xi \pi)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	ADAMOVICH 98	WA89	$\Sigma^-$ nucleus, 345 GeV/c	

$\Gamma(\Xi^-\pi^+\pi^0)/\Gamma(\Sigma\bar{K})$						$\Gamma_4/\Gamma_2$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT		
<0.04	DIONISI	78	HBC	0	$K^-p$ 4.2 GeV/c	

$\Gamma(\Xi^-\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$					
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
possibly seen	4	BIAGI	87	SPEC	—	$\Xi^-$ Be 116 GeV

$\Gamma(\Xi^-\pi^+\pi^-)/\Gamma(\Sigma\bar{K})$						$\Gamma_5/\Gamma_2$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT		
<0.03	DIONISI	78	HBC	—	$K^-p$ 4.2 GeV/c	

$\Gamma(\Xi(1530)\pi)/\Gamma(\Sigma K)$	$\Gamma_6/\Gamma_2$				
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<0.06	DIONISI	78	HBC	—	$K^-p$ 4.2 GeV/c

### $\Xi(1690)$ FOOTNOTES

- <sup>1</sup> From a fit to the  $\Sigma^+ K^-$  spectrum.
- <sup>2</sup> From a coupled-channel analysis of the  $\Sigma^+ K^-$  and  $\Lambda \bar{K}^0$  spectra.
- <sup>3</sup> A fit to the inclusive spectrum from  $\Xi^- N \rightarrow \Lambda K^- X$ .
- <sup>4</sup> From a coupled-channel analysis of the  $\Sigma^0 K^-$  and  $\Lambda K^-$  spectra.

### $\Xi(1690)$ REFERENCES

AUBERT	08AK PR D78 034008	B. Aubert <i>et al.</i>	(BABAR Collab.)
ABE	02C PL B524 33	K. Abe <i>et al.</i>	(KEK BELLE Collab.)
ADAMOVICH	98 EPJ C5 621	M.I. Adamovich <i>et al.</i>	(CERN WA89 Collab.)
BIAGI	87 ZPHY C34 15	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
BIAGI	81 ZPHY C9 305	S.F. Biagi <i>et al.</i>	(BRIS, CAVE, GEVA+)
DIONISI	78 PL 80B 145	C. Dionisi <i>et al.</i>	(CERN, AMST, NUM+)

## $\Xi(1820) 3/2^-$

$$I(J^P) = \frac{1}{2}(\frac{3}{2}^-) \text{ Status: } ***$$

The clearest evidence is an 8-standard-deviation peak in  $\Lambda K^-$  seen by GAY 76C. TEODORO 78 favors  $J = 3/2$ , but cannot make a parity discrimination. BIAGI 87C is consistent with  $J = 3/2$  and favors negative parity for this  $J$  value.

### $\Xi(1820)$ MASS

We only average the measurements that appear to us to be most significant and best determined.

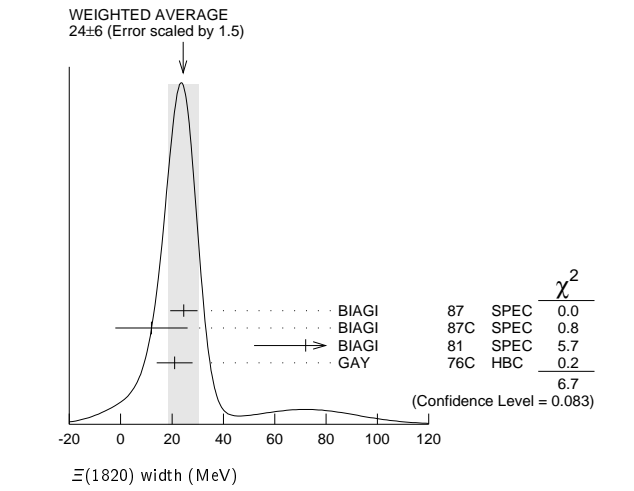
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1823 ± 5 OUR ESTIMATE</b>					
<b>1823.4 ± 1.4 OUR AVERAGE</b>					
1819.4 ± 3.1 ± 2.0	280	<sup>1</sup> BIAGI	87 SPEC	0	$\Xi^-$ Be → ( $\Lambda K^-$ ) X
1826 ± 3 ± 1	54	BIAGI	87c SPEC	0	$\Xi^-$ Be → ( $\Lambda \bar{K}^0$ ) X
1822 ± 6		JENKINS	83 MPS	—	$K^- p \rightarrow K^+ (\text{MM})$
1830 ± 6	300	BIAGI	81 SPEC	—	SPS hyperon beam
1823 ± 2	130	GAY	76c HBC	—	$K^- p$ 4.2 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1817 ± 3		ADAMOVICH	99B WA89		$\Sigma^-$ nucleus, 345 GeV
1797 ± 19	74	BRIEFEL	77 HBC	0	$K^- p$ 2.87 GeV/c
1829 ± 9	68	BRIEFEL	77 HBC	—0	$\Xi(1530) \pi$
1860 ± 14	39	BRIEFEL	77 HBC	—	$\Sigma^- \bar{K}^0$
1870 ± 9	44	BRIEFEL	77 HBC	0	$\Lambda \bar{K}^0$
1813 ± 4	57	BRIEFEL	77 HBC	—	$\Lambda K^-$
1807 ± 27		DIBIANCA	75 DBC	—0	$\Xi \pi \pi, \Xi^* \pi$
1762 ± 8	28	<sup>2</sup> BADIAR	72 HBC	—0	$\Xi \pi, \Xi \pi \pi, Y K$
1838 ± 5	38	<sup>2</sup> BADIAR	72 HBC	—0	$\Xi \pi, \Xi \pi \pi, Y K$
1830 ± 10	25	<sup>3</sup> CRENNELL	70B DBC	—0	3.6, 3.9 GeV/c
1826 ± 12		<sup>4</sup> CRENNELL	70B DBC	—0	3.6, 3.9 GeV/c
1830 ± 10	40	ALITTI	69 HBC	—	$\Lambda, \Sigma \bar{K}$
1814 ± 4	30	BADIAR	65 HBC	0	$\Lambda \bar{K}^0$
1817 ± 7	29	SMITH	65c HBC	—0	$\Lambda \bar{K}^0, \Lambda K^-$
1770		HALSTEINSLID63	FBC	—0	$K^-$ freon 3.5 GeV/c

### $\Xi(1820)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>24 +15 -10 OUR ESTIMATE</b>					
<b>24 ± 6 OUR AVERAGE</b>					Error includes scale factor of 1.5. See the ideogram below.
24.6 ± 5.3	280	<sup>1</sup> BIAGI	87 SPEC	0	$\Xi^-$ Be → ( $\Lambda K^-$ ) X
12 ± 14 ± 1.7	54	BIAGI	87c SPEC	0	$\Xi^-$ Be → ( $\Lambda \bar{K}^0$ ) X
72 ± 20	300	BIAGI	81 SPEC	—	SPS hyperon beam
21 ± 7	130	GAY	76c HBC	—	$K^- p$ 4.2 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •					
23 ± 13		ADAMOVICH	99B WA89		$\Sigma^-$ nucleus, 345 GeV
99 ± 57	74	BRIEFEL	77 HBC	0	$K^- p$ 2.87 GeV/c
52 ± 34	68	BRIEFEL	77 HBC	—0	$\Xi(1530) \pi$
72 ± 17	39	BRIEFEL	77 HBC	—	$\Sigma^- \bar{K}^0$
44 ± 11	44	BRIEFEL	77 HBC	0	$\Lambda \bar{K}^0$
26 ± 11	57	BRIEFEL	77 HBC	—	$\Lambda K^-$
85 ± 58		DIBIANCA	75 DBC	—0	$\Xi \pi \pi, \Xi^* \pi$
51 ± 13		<sup>2</sup> BADIAR	72 HBC	—0	Lower mass
58 ± 13		<sup>2</sup> BADIAR	72 HBC	—0	Higher mass
103 +38 -24		<sup>3</sup> CRENNELL	70B DBC	—0	3.6, 3.9 GeV/c
48 +36 -19		<sup>4</sup> CRENNELL	70B DBC	—0	3.6, 3.9 GeV/c
55 +40 -20		ALITTI	69 HBC	—	$\Lambda, \Sigma \bar{K}$
12 ± 4		BADIAR	65 HBC	0	$\Lambda \bar{K}^0$
30 ± 7		SMITH	65B HBC	—0	$\Lambda \bar{K}$
<80		HALSTEINSLID63	FBC	—0	$K^-$ freon 3.5 GeV/c

# Baryon Particle Listings

## $\Xi(1820), \Xi(1950)$



### $\Xi(1820)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda \bar{K}$	large
$\Gamma_2 \Sigma \bar{K}$	small
$\Gamma_3 \Xi \pi$	small
$\Gamma_4 \Xi(1530)\pi$	small
$\Gamma_5 \Xi \pi \pi$ (not $\Xi(1530)\pi$ )	

### $\Xi(1820)$ BRANCHING RATIOS

The dominant modes seem to be  $\Lambda \bar{K}$  and (perhaps)  $\Xi(1530)\pi$ , but the branching fractions are very poorly determined.

$\Gamma(\Lambda \bar{K})/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.25±0.05 OUR AVERAGE</b>				
0.24±0.05	ANISOVICH	12A	DPWA	Multichannel
0.30±0.15	ALITTI	69	HBC	$K^- p$ 3.9–5 GeV/c

$\Gamma(\Xi \pi)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.10±0.10</b>	ALITTI	69	HBC	$K^- p$ 3.9–5 GeV/c

$\Gamma(\Xi \pi)/\Gamma(\Lambda \bar{K})$				$\Gamma_3/\Gamma_1$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;0.36</b>	95			
0.20±0.20	GAY	76c	HBC	$K^- p$ 4.2 GeV/c
	BADIER	65	HBC	$K^- p$ 3 GeV/c

$\Gamma(\Xi \pi)/\Gamma(\Xi(1530)\pi)$				$\Gamma_3/\Gamma_4$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.5<sup>+0.6</sup><sub>-0.4</sub></b>	APSELL	70	HBC	$K^- p$ 2.87 GeV/c

$\Gamma(\Sigma \bar{K})/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.30±0.15</b>	ALITTI	69	HBC	$K^- p$ 3.9–5 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.02	TRIPP	67	RVUE	Use SMITH 65c

$\Gamma(\Sigma \bar{K})/\Gamma(\Lambda \bar{K})$				$\Gamma_2/\Gamma_1$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.24±0.10</b>	GAY	76c	HBC	$K^- p$ 4.2 GeV/c

$\Gamma(\Xi(1530)\pi)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.30±0.15</b>	ALITTI	69	HBC	$K^- p$ 3.9–5 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	ASTON	85B	LASS	$K^- p$ 11 GeV/c
not seen	<sup>5</sup> HASSALL	81	HBC	$K^- p$ 6.5 GeV/c
<0.25	<sup>6</sup> DAUBER	69	HBC	$K^- p$ 2.7 GeV/c

$\Gamma(\Xi(1530)\pi)/\Gamma(\Lambda \bar{K})$				$\Gamma_4/\Gamma_1$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.38±0.27 OUR AVERAGE</b>				Error includes scale factor of 2.3.
1.0 ±0.3	GAY	76c	HBC	$K^- p$ 4.2 GeV/c
0.26±0.13	SMITH	65c	HBC	$K^- p$ 2.45–2.7 GeV/c

$\Gamma(\Xi \pi \pi$ (not $\Xi(1530)\pi$ ))/ $\Gamma(\Lambda \bar{K})$				$\Gamma_5/\Gamma_1$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.30±0.20</b>	BIAGI	87	SPEC	$\Xi^- \text{Be}$ 116 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.14	<sup>7</sup> BADIER	65	HBC	0 1 st. dev. limit
>0.1	SMITH	65c	HBC	$K^- p$ 2.45–2.7 GeV/c

$\Gamma(\Xi \pi \pi$ (not $\Xi(1530)\pi$ ))/ $\Gamma(\Xi(1530)\pi)$				$\Gamma_5/\Gamma_4$
VALUE	DOCUMENT ID	TECN	CHG	COMMENT
consistent with zero	GAY	76c	HBC	$K^- p$ 4.2 GeV/c
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.3±0.5	<sup>8</sup> APSELL	70	HBC	$K^- p$ 2.87 GeV/c

### $\Xi(1820)$ FOOTNOTES

- <sup>1</sup> BIAGI 87 also sees weak signals in the in the  $\Xi^- \pi^+ \pi^-$  channel at 1782.6 ± 1.4 MeV ( $\Gamma = 6.0 \pm 1.5$  MeV) and 1831.9 ± 2.8 MeV ( $\Gamma = 9.6 \pm 9.9$  MeV).
- <sup>2</sup> BADIER 72 adds all channels and divides the peak into lower and higher mass regions. The data can also be fitted with a single Breit-Wigner of mass 1800 MeV and width 150 MeV.
- <sup>3</sup> From a fit to inclusive  $\Xi \pi$ ,  $\Xi \pi \pi$ , and  $\Lambda K^-$  spectra.
- <sup>4</sup> From a fit to inclusive  $\Xi \pi$  and  $\Xi \pi \pi$  spectra only.
- <sup>5</sup> Including  $\Xi \pi \pi$ .
- <sup>6</sup> DAUBER 69 uses in part the same data as SMITH 65c.
- <sup>7</sup> For the decay mode  $\Xi^- \pi^+ \pi^0$  only. This limit includes  $\Xi(1530)\pi$ .
- <sup>8</sup> Or less. Upper limit for the 3-body decay.

### $\Xi(1820)$ REFERENCES

ANISOVICH	12A	EPJ A48 15	A.V. Anisovich <i>et al.</i>	(BONN, PNPI)
ADAMOVICH	99B	EPJ C11 271	M.I. Adamovich <i>et al.</i>	(CERN WA89 Collab.)
BIAGI	87	ZPHY C34 15	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
BIAGI	87C	ZPHY C34 175	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
ASTON	85B	PR D32 2270	D. Aston <i>et al.</i>	(SLAC, CARL, CNRC, CINC)
JENKINS	83	PRL 51 951	C.M. Jenkins <i>et al.</i>	(FSU, BRAN, LBL+)
BIAGI	81	ZPHY C9 305	S.F. Biagi <i>et al.</i>	(BRIS, CAVE, GEVA+)
HASSALL	81	NP B189 397	J.K. Hassall <i>et al.</i>	(CAVE, MSU)
TEODORO	78	PL 77B 451	D. Teodoro <i>et al.</i>	(AMST, CERN, NUM+)
BRIEFEL	77	PR D16 2706	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
Also		PRL 23 884	S.P. Apseil <i>et al.</i>	(BRAN, UMD, SYRA+)
GAY	76c	PL 62B 477	J.B. Gay <i>et al.</i>	(AMST, CERN, NUM)IJ
DIBIANCA	75	NP B98 137	F.A. Dibianca, R.J. Endorf	(CMU)
BADIER	72	NP B37 429	J. Badier <i>et al.</i>	(EPOL)
APSELL	70	PRL 24 777	S.P. Apseil <i>et al.</i>	(BRAN, UMD, SYRA+)
CRENNELL	70B	PR D1 847	D.J. Crennell <i>et al.</i>	(BNL)
ALITTI	69	PR 22 79	J. Alitti <i>et al.</i>	(BNL, SYRA)I
DAUBER	69	PR 179 1262	P.M. Dauber <i>et al.</i>	(LRL)
TRIPP	67	NP B3 10	R.D. Tripp <i>et al.</i>	(LRL, SLAC, CERN+)
BADIER	65	PL 16 171	J. Badier <i>et al.</i>	(EPOL, SACL, AMST)I
SMITH	65B	Athens Conf. 251	G.A. Smith, J.S. Lindsey	(LRL)
SMITH	65C	PRL 14 25	G.A. Smith <i>et al.</i>	(LRL)IJP
HALSTEINSLID	63	Siena Conf. 1 73	A. Halsteinslid <i>et al.</i>	(BERG, CERN, EPOL+)

### OTHER RELATED PAPERS

TEODORO	78	PL 77B 451	D. Teodoro <i>et al.</i>	(AMST, CERN, NUM+)
BRIEFEL	75	PR D12 1859	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
SCHMIDT	73	Purdue Conf. 363	P.E. Schmidt	(BRAN)
MERRILL	68	PR 167 1202	D.W. Merrill, J. Button-Shafer	(LRL)
SMITH	64	PRL 13 61	G.A. Smith <i>et al.</i>	(LRL)IJP

## $\Xi(1950)$

$I(J^P) = \frac{1}{2}(?)^?$  Status: \* \* \*

We list here everything reported between 1875 and 2000 MeV. The accumulated evidence for a  $\Xi$  near 1950 MeV seems strong enough to include a  $\Xi(1950)$  in the main Baryon Table, but not much can be said about its properties. In fact, there may be more than one  $\Xi$  near this mass.

### $\Xi(1950)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1950±15 OUR ESTIMATE</b>				
1955 ± 6		ADAMOVICH	99B	WA89 $\Sigma^-$ nucleus, 345 GeV
1944 ± 9	129	BIAGI	87	SPEC $\Xi^- \text{Be} \rightarrow (\Xi^- \pi^+) \pi^- X$
1963 ± 5 ± 2	63	BIAGI	87c	SPEC $\Xi^- \text{Be} \rightarrow (\Lambda \bar{K}^0) X$
1937 ± 7	150	BIAGI	81	SPEC SPS hyperon beam
1961 ± 18	139	BRIEFEL	77	HBC $2.87 K^- p \rightarrow \Xi^- \pi^+ X$
1936 ± 22	44	BRIEFEL	77	HBC $2.87 K^- p \rightarrow \Xi^0 \pi^- X$
1964 ± 10	56	BRIEFEL	77	HBC $\Xi(1530)\pi$
1900 ± 12		DIBIANCA	75	DBC $\Xi \pi$
1952 ± 11	25	ROSS	73c	( $\Xi \pi$ ) $^-$
1956 ± 6	29	BADIER	72	HBC $\Xi \pi, \Xi \pi \pi, Y K$
1955 ± 14	21	GOLDWASSER	70	HBC $\Xi \pi$
1894 ± 18	66	DAUBER	69	HBC $\Xi \pi$
1930 ± 20	27	ALITTI	68	HBC $\Xi^- \pi^+$
1933 ± 16	35	BADIER	65	HBC $\Xi^- \pi^+$

See key on page 601

## Baryon Particle Listings

 $\Xi(1950), \Xi(2030)$  $\Xi(1950)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>60±20 OUR ESTIMATE</b>				
68±22		ADAMOVICH 99B	WA89	$\Sigma^-$ nucleus, 345 GeV
100±31	129	BIAGI 87	SPEC	$\Xi^- \text{Be} \rightarrow (\Xi^- \pi^+) \pi^- X$
25±15±1.2	63	BIAGI 87c	SPEC	$\Xi^- \text{Be} \rightarrow (\Lambda \bar{K}^0) X$
60±8	150	BIAGI 81	SPEC	SPS hyperon beam
159±57	139	BRIEFEL 77	HBC	$2.87 K^- p \rightarrow \Xi^- \pi^+ X$
87±26	44	BRIEFEL 77	HBC	$2.87 K^- p \rightarrow \Xi^0 \pi^- X$
60±39	56	BRIEFEL 77	HBC	$\Xi(1530) \pi$
63±78		DIBIANCA 75	DBC	$\Xi \pi$
38±10		ROSS 73c		$(\Xi \pi)^-$
35±11	29	BADIER 72	HBC	$\Xi \pi, \Xi \pi \pi, \Upsilon K$
56±26	21	GOLDWASSER 70	HBC	$\Xi \pi$
98±23	66	DAUBER 69	HBC	$\Xi \pi$
80±40	27	ALITTI 68	HBC	$\Xi^- \pi^+$
140±35	35	BADIER 65	HBC	$\Xi^- \pi^+$

 $\Xi(1950)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda \bar{K}$	seen
$\Gamma_2 \Sigma \bar{K}$	possibly seen
$\Gamma_3 \Xi \pi$	seen
$\Gamma_4 \Xi(1530) \pi$	
$\Gamma_5 \Xi \pi \pi$ (not $\Xi(1530) \pi$ )	

 $\Xi(1950)$  BRANCHING RATIOS

$\Gamma(\Sigma \bar{K})/\Gamma(\Lambda \bar{K})$		$\Gamma_2/\Gamma_1$	
VALUE	CL% EVTS	DOCUMENT ID	TECN COMMENT
<2.3	90 0	BIAGI 87c	SPEC $\Xi^- \text{Be}$ 116 GeV
$\Gamma(\Sigma \bar{K})/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
possibly seen	17	HASSALL 81	HBC $K^- p$ 6.5 GeV/c
$\Gamma(\Xi \pi)/\Gamma(\Xi(1530) \pi)$		$\Gamma_3/\Gamma_4$	
VALUE	DOCUMENT ID	TECN	COMMENT
2.8+0.7-0.6	APSELL 70	HBC	
$\Gamma(\Xi \pi \pi \text{ (not } \Xi(1530) \pi))/\Gamma(\Xi(1530) \pi)$		$\Gamma_5/\Gamma_4$	
VALUE	DOCUMENT ID	TECN	COMMENT
0.0±0.3	APSELL 70	HBC	

 $\Xi(1950)$  REFERENCES

ADAMOVICH 99B	EPJ C11 271	M.I. Adamovich <i>et al.</i>	(CERN WA89 Collab.)
BIAGI 87	ZPHY C34 15	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
BIAGI 87c	ZPHY C34 175	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
BIAGI 81	ZPHY C9 305	S.F. Biagi <i>et al.</i>	(BRIS, CAVE, GEVA+)
HASSALL 81	NP B189 397	J.K. Hassall <i>et al.</i>	(CAVE, MSU)
BRIEFEL 77	PR D16 2706	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
Also	Duke Conf. 317	E. Briefel <i>et al.</i>	(BRAN, UMD, SYRA+)
Hyperon Resonances, 1970			
DIBIANCA 75	NP B98 137	F.A. Dibianca, R.J. Endorf	(CMU)
ROSS 73c	Purdue Conf. 345	R.T. Ross, J.L. Lloyd, D. Radojicic	(OXF)
BADIER 72	NP B37 429	J. Badier <i>et al.</i>	(EPOL)
APSELL 70	PRL 24 777	S.P. Appell <i>et al.</i>	(BRAN, UMD, SYRA+)
GOLDWASSER 70	PR D1 1960	E.L. Goldwasser, P.F. Schultz	(ILL)
DAUBER 69	PR 179 1262	P.M. Dauber <i>et al.</i>	(LRL)
ALITTI 68	PRL 21 1119	J. Alitti <i>et al.</i>	(BNL, SYRA)
BADIER 65	PL 16 171	J. Badier <i>et al.</i>	(EPOL, SAFL, AMST)

 $\Xi(2030)$ 

$$J(P) = \frac{1}{2} (\geq \frac{5}{2}?) \text{ status: } ***$$

The evidence for this state has been much improved by HEMINGWAY 77, who see an eight standard deviation enhancement in  $\Sigma \bar{K}$  and a weaker coupling to  $\Lambda \bar{K}$ . ALITTI 68 and HEMINGWAY 77 observe no signals in the  $\Xi \pi \pi$  (or  $\Xi(1530) \pi$ ) channel, in contrast to DIBIANCA 75. The decay  $(\Lambda/\Sigma) \bar{K} \pi$  reported by BARTSCH 69 is also not confirmed by HEMINGWAY 77.

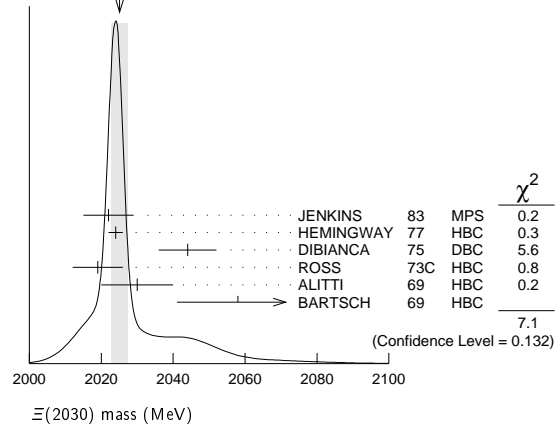
A moments analysis of the HEMINGWAY 77 data indicates at a level of three standard deviations that  $J \geq 5/2$ .

 $\Xi(2030)$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2025 ± 5 OUR ESTIMATE</b>					
<b>2025.1±2.4 OUR AVERAGE</b> Error includes scale factor of 1.3. See the ideogram below.					
2022 ± 7		JENKINS 83	MPS	—	$K^- p \rightarrow K^+ \text{MM}$
2024 ± 2	200	HEMINGWAY 77	HBC	—	$K^- p$ 4.2 GeV/c
2044 ± 8		DIBIANCA 75	DBC	—0	$\Xi \pi \pi, \Xi^* \pi$

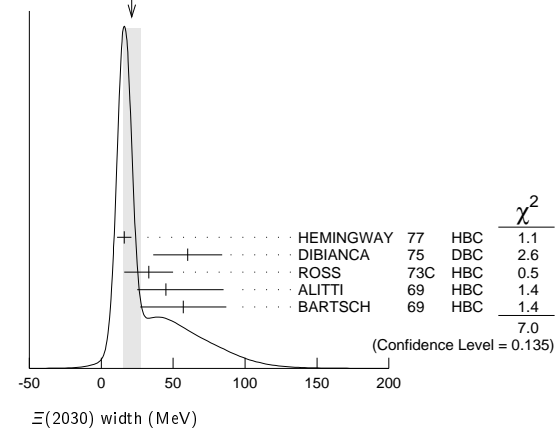
2019 ± 7	15	ROSS 73c	HBC	—0	$\Sigma \bar{K}$
2030 ± 10	42	ALITTI 69	HBC	—	$K^- p$ 3.9–5 GeV/c
2058 ± 17	40	BARTSCH 69	HBC	—0	$K^- p$ 10 GeV/c

WEIGHTED AVERAGE  
2025.1±2.4 (Error scaled by 1.3)

 $\Xi(2030)$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>20±15 OUR ESTIMATE</b>					
<b>21±6 OUR AVERAGE</b> Error includes scale factor of 1.3. See the ideogram below.					
16±5	200	HEMINGWAY 77	HBC	—	$K^- p$ 4.2 GeV/c
60±24		DIBIANCA 75	DBC	—0	$\Xi \pi \pi, \Xi^* \pi$
33±17	15	ROSS 73c	HBC	—0	$\Sigma \bar{K}$
45+40-20		ALITTI 69	HBC	—	$K^- p$ 3.9–5 GeV/c
57±30		BARTSCH 69	HBC	—0	$K^- p$ 10 GeV/c

WEIGHTED AVERAGE  
21±6 (Error scaled by 1.3)

 $\Xi(2030)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda \bar{K}$	~ 20 %
$\Gamma_2 \Sigma \bar{K}$	~ 80 %
$\Gamma_3 \Xi \pi$	small
$\Gamma_4 \Xi(1530) \pi$	small
$\Gamma_5 \Xi \pi \pi$ (not $\Xi(1530) \pi$ )	small
$\Gamma_6 \Lambda \bar{K} \pi$	small
$\Gamma_7 \Sigma \bar{K} \pi$	small

 $\Xi(2030)$  BRANCHING RATIOS

$\Gamma(\Xi \pi)/[\Gamma(\Lambda \bar{K}) + \Gamma(\Sigma \bar{K}) + \Gamma(\Xi \pi) + \Gamma(\Xi(1530) \pi)]$		$\Gamma_3/(\Gamma_1 + \Gamma_2 + \Gamma_3 + \Gamma_4)$	
VALUE	DOCUMENT ID	TECN	CHG COMMENT
••• We do not use the following data for averages, fits, limits, etc. •••			
<0.30	ALITTI 69	HBC	— 1 standard dev. limit

# Baryon Particle Listings

## $\Xi(2030)$ , $\Xi(2120)$ , $\Xi(2250)$ , $\Xi(2370)$

$\Gamma(\Xi\pi)/\Gamma(\Sigma\bar{K})$					$\Gamma_3/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.19	95	HEMINGWAY 77	HBC	—	$K^-p$ 4.2 GeV/c	
$\Gamma(\Lambda\bar{K})/[\Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi\pi) + \Gamma(\Xi(1530)\pi)]$					$\Gamma_1/(\Gamma_1+\Gamma_2+\Gamma_3+\Gamma_4)$	
VALUE		DOCUMENT ID	TECN	CHG	COMMENT	
0.25±0.15		ALITTI 69	HBC	—	$K^-p$ 3.9-5 GeV/c	
$\Gamma(\Lambda\bar{K})/\Gamma(\Sigma\bar{K})$					$\Gamma_1/\Gamma_2$	
VALUE		DOCUMENT ID	TECN	CHG	COMMENT	
0.22±0.09		HEMINGWAY 77	HBC	—	$K^-p$ 4.2 GeV/c	
$\Gamma(\Sigma\bar{K})/[\Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi\pi) + \Gamma(\Xi(1530)\pi)]$					$\Gamma_2/(\Gamma_1+\Gamma_2+\Gamma_3+\Gamma_4)$	
VALUE		DOCUMENT ID	TECN	CHG	COMMENT	
0.75±0.20		ALITTI 69	HBC	—	$K^-p$ 3.9-5 GeV/c	
$\Gamma(\Xi(1530)\pi)/[\Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi\pi) + \Gamma(\Xi(1530)\pi)]$					$\Gamma_4/(\Gamma_1+\Gamma_2+\Gamma_3+\Gamma_4)$	
VALUE		DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.15		ALITTI 69	HBC	—	1 standard dev. limit	
$[\Gamma(\Xi(1530)\pi) + \Gamma(\Xi\pi\pi(\text{not } \Xi(1530)\pi))]/\Gamma(\Sigma\bar{K})$					$(\Gamma_4+\Gamma_5)/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.11	95	<sup>1</sup> HEMINGWAY 77	HBC	—	$K^-p$ 4.2 GeV/c	
$\Gamma(\Lambda\bar{K}\pi)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$	
VALUE		DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
seen		BARTSCH 69	HBC	$K^-p$ 10 GeV		
$\Gamma(\Lambda\bar{K}\pi)/\Gamma(\Sigma\bar{K})$					$\Gamma_6/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.32	95	HEMINGWAY 77	HBC	—	$K^-p$ 4.2 GeV/c	
$\Gamma(\Sigma\bar{K}\pi)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$	
VALUE		DOCUMENT ID	TECN	COMMENT		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
seen		BARTSCH 69	HBC	$K^-p$ 10 GeV		
$\Gamma(\Sigma\bar{K}\pi)/\Gamma(\Sigma\bar{K})$					$\Gamma_7/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT	
<0.04	95	<sup>2</sup> HEMINGWAY 77	HBC	—	$K^-p$ 4.2 GeV/c	

### $\Xi(2030)$ FOOTNOTES

- <sup>1</sup> For the decay mode  $\Xi^- \pi^+ \pi^-$  only.  
<sup>2</sup> For the decay mode  $\Sigma^\pm K^- \pi^\mp$  only.

### $\Xi(2030)$ REFERENCES

JENKINS 83	PRL 51 951	C.M. Jenkins <i>et al.</i>	(FSU, BRAN, LBL+)
HEMINGWAY 77	PL 68B 197	R.J. Hemingway <i>et al.</i>	(AMST, CERN, NIJM+)
Also	PL 62B 477	J.B. Gay <i>et al.</i>	(AMST, CERN, NIJM)
DIBIANCA 75	NP B98 137	F.A. Dibianca, R.J. Endorf	(CMU)
ROSS 73C	Purdue Conf. 345	R.T. Ross, J.L. Lloyd, D. Radojicic	(OXF)
ALITTI 69	PRL 22 79	J. Alitti <i>et al.</i>	(BNL, SYRA)
BARTSCH 69	PL 28B 439	J. Bartsch <i>et al.</i>	(AACH, BERL, CERN+)
ALITTI 68	PRL 21 1119	J. Alitti <i>et al.</i>	(BNL, SYRA)

$\Xi(2120)$	$I(J^P) = \frac{1}{2}(?)^?$ Status: * <i>J, P</i> need confirmation.
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OMITTED FROM SUMMARY TABLE

### $\Xi(2120)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$\approx 2120$ OUR ESTIMATE				
2137±4	18	<sup>1</sup> CHLIAPNIK... 79	HBC	$K^+p$ 32 GeV/c
2123±7		<sup>2</sup> GAY 76c	HBC	$K^-p$ 4.2 GeV/c

### $\Xi(2120)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<20	18	<sup>1</sup> CHLIAPNIK... 79	HBC	$K^+p$ 32 GeV/c
25±12		<sup>2</sup> GAY 76c	HBC	$K^-p$ 4.2 GeV/c

### $\Xi(2120)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\Lambda\bar{K}$	seen

### $\Xi(2120)$ BRANCHING RATIOS

$\Gamma(\Lambda\bar{K})/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT	$\Gamma_1/\Gamma$
VALUE				
seen	<sup>1</sup> CHLIAPNIK... 79	HBC	$K^+p \rightarrow (\bar{\Lambda} K^+) X$	
seen	<sup>2</sup> GAY 76c	HBC	$K^-p$ 4.2 GeV/c	

### $\Xi(2120)$ FOOTNOTES

- <sup>1</sup> CHLIAPNIKOV 79 does not uniquely identify the  $K^+$  in the  $(\bar{\Lambda} K^+) X$  final state. It also reports bumps with fewer events at 2240, 2540, and 2830 MeV.  
<sup>2</sup> GAY 76c sees a 4-standard deviation signal. However, HEMINGWAY 77, with more events from the same experiment points out that the signal is greatly reduced if a cut is made on the 4-momentum  $u$ . This suggests an anomalous production mechanism if the  $\Xi(2120)$  is real.

### $\Xi(2120)$ REFERENCES

CHLIAPNIK... 79	NP B158 253	P.V. Chliapnikov <i>et al.</i>	(CERN, BELG, MONS)
HEMINGWAY 77	PL 68B 197	R.J. Hemingway <i>et al.</i>	(AMST, CERN, NIJM+)
GAY 76c	PL 62B 477	J.B. Gay <i>et al.</i>	(AMST, CERN, NIJM)

## $\Xi(2250)$

$I(J^P) = \frac{1}{2}(?)^?$  Status: \* \*  
*J, P* need confirmation.

OMITTED FROM SUMMARY TABLE

The evidence for this state is mixed. BARTSCH 69 sees a bump of not much statistical significance in  $\Lambda\bar{K}\pi$ ,  $\Sigma\bar{K}\pi$ , and  $\Xi\pi\pi$  mass spectra. GOLDWASSER 70 sees a narrower bump in  $\Xi\pi\pi$  at a higher mass. Not seen by HASSALL 81 with 45 events/ $\mu\text{b}$  at 6.5 GeV/c. Seen by JENKINS 83. Perhaps seen by BIAGI 87.

### $\Xi(2250)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$\approx 2250$ OUR ESTIMATE					
2189± 7	66	BIAGI 87	SPEC	—	$\Xi^- \text{Be} \rightarrow (\Xi^- \pi^+ \pi^-) X$
2214± 5		JENKINS 83	MPS	—	$K^-p \rightarrow K^+ \text{MM}$
2295±15	18	GOLDWASSER 70	HBC	—	$K^-p$ 5.5 GeV/c
2244±52	35	BARTSCH 69	HBC		$K^-p$ 10 GeV/c

### $\Xi(2250)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
46±27	66	BIAGI 87	SPEC	—	$\Xi^- \text{Be} \rightarrow (\Xi^- \pi^+ \pi^-) X$
< 30		GOLDWASSER 70	HBC	—	$K^-p$ 5.5 GeV/c
130±80		BARTSCH 69	HBC		

### $\Xi(2250)$ DECAY MODES

Mode
$\Gamma_1$ $\Xi\pi\pi$
$\Gamma_2$ $\Lambda\bar{K}\pi$
$\Gamma_3$ $\Sigma\bar{K}\pi$

### $\Xi(2250)$ REFERENCES

BIAGI 87	ZPHY C34 15	S.F. Biagi <i>et al.</i>	(BRIS, CERN, GEVA+)
JENKINS 83	PRL 51 951	C.M. Jenkins <i>et al.</i>	(FSU, BRAN, LBL+)
HASSALL 81	NP B189 397	J.K. Hassall <i>et al.</i>	(CAVE, MSU)
GOLDWASSER 70	PR D1 1960	E.L. Goldwasser, P.F. Schultz	(ILL)
BARTSCH 69	PL 28B 439	J. Bartsch <i>et al.</i>	(AACH, BERL, CERN+)

## $\Xi(2370)$

$I(J^P) = \frac{1}{2}(?)^?$  Status: \* \*  
*J, P* need confirmation.

OMITTED FROM SUMMARY TABLE

### $\Xi(2370)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$\approx 2370$ OUR ESTIMATE					
2356±10		JENKINS 83	MPS	—	$K^-p \rightarrow K^+ \text{MM}$
2370	50	HASSALL 81	HBC	—0	$K^-p$ 6.5 GeV/c
2373± 8	94	AMIRZADEH 80	HBC	—0	$K^-p$ 8.25 GeV/c
2392±27		DIBIANCA 75	DBC		$\Xi 2\pi$

### $\Xi(2370)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
80	50	HASSALL 81	HBC	—0	$K^-p$ 6.5 GeV/c
80±25	94	AMIRZADEH 80	HBC	—0	$K^-p$ 8.25 GeV/c
75±69		DIBIANCA 75	DBC		$\Xi 2\pi$

See key on page 601

Baryon Particle Listings  
 $\Xi(2370), \Xi(2500)$

$\Xi(2370)$ DECAY MODES					
Mode		Fraction ( $\Gamma_i/\Gamma$ )			
$\Gamma_1$	$\Lambda\bar{K}\pi$ Includes $\Gamma_4 + \Gamma_6$ .	seen			
$\Gamma_2$	$\Sigma\bar{K}\pi$ Includes $\Gamma_5 + \Gamma_6$ .	seen			
$\Gamma_3$	$\Omega^-K$				
$\Gamma_4$	$\Lambda\bar{K}^*(892)$				
$\Gamma_5$	$\Sigma\bar{K}^*(892)$				
$\Gamma_6$	$\Sigma(1385)\bar{K}$				
$\Xi(2370)$ BRANCHING RATIOS					
$\Gamma(\Lambda\bar{K}\pi)/\Gamma_{\text{total}}$		$\Gamma_1/\Gamma$			
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
seen		AMIRZADEH 80	HBC	−0	$K^-p$ 8.25 GeV/c
$\Gamma(\Sigma\bar{K}\pi)/\Gamma_{\text{total}}$		$\Gamma_2/\Gamma$			
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
seen		AMIRZADEH 80	HBC	−0	$K^-p$ 8.25 GeV/c
$[\Gamma(\Lambda\bar{K}\pi) + \Gamma(\Sigma\bar{K}\pi)]/\Gamma_{\text{total}}$		$(\Gamma_1+\Gamma_2)/\Gamma$			
VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
seen	50	HASSALL 81	HBC	−0	$K^-p$ 6.5 GeV/c
$\Gamma(\Omega^-K)/\Gamma_{\text{total}}$		$\Gamma_3/\Gamma$			
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
0.09±0.04		<sup>1</sup> KINSON 80	HBC	—	$K^-p$ 8.25 GeV/c
$[\Gamma(\Lambda\bar{K}^*(892)) + \Gamma(\Sigma\bar{K}^*(892))]/\Gamma_{\text{total}}$		$(\Gamma_4+\Gamma_5)/\Gamma$			
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
0.22±0.13		<sup>1</sup> KINSON 80	HBC	—	$K^-p$ 8.25 GeV/c
$\Gamma(\Sigma(1385)\bar{K})/\Gamma_{\text{total}}$		$\Gamma_6/\Gamma$			
VALUE		DOCUMENT ID	TECN	CHG	COMMENT
0.12±0.08		<sup>1</sup> KINSON 80	HBC	—	$K^-p$ 8.25 GeV/c
$\Xi(2370)$ FOOTNOTES					
<sup>1</sup> KINSON 80 is a reanalysis of AMIRZADEH 80 with 50% more events.					

$\Xi(2370)$ REFERENCES					
JENKINS	83	PRL 51 951	C.M. Jenkins <i>et al.</i>	(FSU, BRAN, LBL+)	
HASSALL	81	NP B189 397	J.K. Hassall <i>et al.</i>	(CAVE, MSU)	
AMIRZADEH	80	PL 90B 324	J. Amirzadeh <i>et al.</i>	(BIRM, CERN, GLAS+)	
KINSON	80	Toronto Conf. 263	J.B. Kinson <i>et al.</i>	(BIRM, CERN, GLAS+)	
DIBIANCA	75	NP B98 137	F.A. Dibianca, R.J. Endorf	(CMU)	

$\Xi(2500)$

$I(J^P) = \frac{1}{2}(?)^?$  Status: \*  
 $J, P$  need confirmation.

OMITTED FROM SUMMARY TABLE

The ALITTI 69 peak might be instead the  $\Xi(2370)$  or might be neither the  $\Xi(2370)$  nor the  $\Xi(2500)$ .

$\Xi(2500)$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$\approx$ 2500 OUR ESTIMATE					
25 05 ±10		JENKINS	83	MP S	— $K^-p \rightarrow K^+$
24 30 ±20	30	ALITTI	69	HBC	— $K^-p$ 4.6–5 GeV/c
25 00 ±10	45	BARTSCH	69	HBC	−0 $K^-p$ 10 GeV/c
$\Xi(2500)$ WIDTH					
VALUE (MeV)	DOCUMENT ID	TECN	CHG		
150 +60 −40	ALITTI 69	HBC	—		
59 ±27	BARTSCH 69	HBC	−0		

$\Xi(2500)$ DECAY MODES					
Mode		Fraction ( $\Gamma_i/\Gamma$ )			
$\Gamma_1$	$\Xi\pi$				
$\Gamma_2$	$\Lambda\bar{K}$				
$\Gamma_3$	$\Sigma\bar{K}$				
$\Gamma_4$	$\Xi\pi\pi$	seen			
$\Gamma_5$	$\Xi(1530)\pi$				
$\Gamma_6$	$\Lambda\bar{K}\pi + \Sigma\bar{K}\pi$	seen			
$\Xi(2500)$ BRANCHING RATIOS					
$\Gamma(\Xi\pi)/[\Gamma(\Xi\pi) + \Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi(1530)\pi)]$		$\Gamma_1/(\Gamma_1 + \Gamma_2 + \Gamma_3 + \Gamma_5)$			
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<0.5	ALITTI 69	HBC		1 standard dev. limit	
$\Gamma(\Lambda\bar{K})/[\Gamma(\Xi\pi) + \Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi(1530)\pi)]$		$\Gamma_2/(\Gamma_1 + \Gamma_2 + \Gamma_3 + \Gamma_5)$			
VALUE	DOCUMENT ID	TECN	CHG		
0.5 ± 0.2	ALITTI 69	HBC		—	
$\Gamma(\Sigma\bar{K})/[\Gamma(\Xi\pi) + \Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi(1530)\pi)]$		$\Gamma_3/(\Gamma_1 + \Gamma_2 + \Gamma_3 + \Gamma_5)$			
VALUE	DOCUMENT ID	TECN	CHG		
0.5 ± 0.2	ALITTI 69	HBC		—	
$\Gamma(\Xi(1530)\pi)/[\Gamma(\Xi\pi) + \Gamma(\Lambda\bar{K}) + \Gamma(\Sigma\bar{K}) + \Gamma(\Xi(1530)\pi)]$		$\Gamma_5/(\Gamma_1 + \Gamma_2 + \Gamma_3 + \Gamma_5)$			
VALUE	DOCUMENT ID	TECN	CHG	COMMENT	
<0.2	ALITTI 69	HBC		1 standard dev. limit	
$\Gamma(\Xi\pi\pi)/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$	
VALUE	DOCUMENT ID	TECN	CHG		
seen	BARTSCH 69	HBC	—0		
$[\Gamma(\Lambda\bar{K}\pi) + \Gamma(\Sigma\bar{K}\pi)]/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$	
VALUE	DOCUMENT ID	TECN	CHG		
seen	BARTSCH 69	HBC	—0		
$\Xi(2500)$ REFERENCES					
JENKINS	83	PRL 51 951	C.M. Jenkins <i>et al.</i>		(FSU, BRAN, LBL+)
ALITTI	69	PRL 22 79	J. Alitti <i>et al.</i>		(BNL, SYRA) I
BARTSCH	69	PL 28B 439	J. Bartsch <i>et al.</i>		(AACH, BERL, CERN+)

Baryon Particle Listings

$\Omega^-$

$\Omega$  BARYONS  
( $S = -3, I = 0$ )

$\Omega^- = sss$

$\Omega^-$

$I(J^P) = 0(\frac{3}{2}^+)$  Status: \*\*\*

The unambiguous discovery in both production and decay was by BARNES 64. The quantum numbers follow from the assignment of the particle to the baryon decuplet. DEUTSCHMANN 78 and BAUBILLIER 78 rule out  $J = 1/2$  and find consistency with  $J = 3/2$ . AUBERT, BE 06 finds from the decay angular distributions of  $\Xi_c^0 \rightarrow \Omega^- K^+$  and  $\Omega_c^0 \rightarrow \Omega^- K^+$  that  $J = 3/2$ ; this depends on the spins of the  $\Xi_c^0$  and  $\Omega_c^0$  being  $J = 1/2$ , their supposed values.

We have omitted some results that have been superseded by later experiments. See our earlier editions.

$\Omega^-$  MASS

The fit assumes the  $\Omega^-$  and  $\overline{\Omega}^+$  masses are the same, and averages them together.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1672.45 ± 0.29 OUR FIT</b>				
<b>1672.43 ± 0.32 OUR AVERAGE</b>				
1673 ± 1	100	HARTOUNI 85	SPEC	80–280 GeV $K_L^0 C$
1673.0 ± 0.8	41	BAUBILLIER 78	HBC	8.25 GeV/c $K^- p$
1671.7 ± 0.6	27	HEMINGWAY 78	HBC	4.2 GeV/c $K^- p$
1673.4 ± 1.7	4	<sup>1</sup> DIBIANCA 75	DBC	4.9 GeV/c $K^- d$
1673.3 ± 1.0	3	PALMER 68	HBC	$K^- p$ 4.6, 5 GeV/c
1671.8 ± 0.8	3	SCHULTZ 68	HBC	$K^- p$ 5.5 GeV/c
1674.2 ± 1.6	5	SCOTTER 68	HBC	$K^- p$ 6 GeV/c
1672.1 ± 1.0	1	<sup>2</sup> FRY 55	EMUL	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1671.43 ± 0.78	13	<sup>3</sup> DEUTSCH... 73	HBC	$K^- p$ 10 GeV/c
1671.9 ± 1.2	6	<sup>3</sup> SPETH 69	HBC	See DEUTSCHMANN 73
1673.0 ± 8.0	1	ABRAMS 64	HBC	$\rightarrow \Xi^- \pi^0$
1670.6 ± 1.0	1	<sup>2</sup> FRY 55B	EMUL	
1615	1	<sup>4</sup> EISENBERG 54	EMUL	

- <sup>1</sup>DIBIANCA 75 gives a mass for each event. We quote the average.
- <sup>2</sup>The FRY 55 and FRY 55B events were identified as  $\Omega^-$  by ALVAREZ 73. The masses assume decay to  $\Lambda K^-$  at rest. For FRY 55B, decay from an atomic orbit could Doppler shift the  $K^-$  energy and the resulting  $\Omega^-$  mass by several MeV. This shift is negligible for FRY 55 because the  $\Omega$  decay is approximately perpendicular to its orbital velocity, as is known because the  $\Lambda$  strikes the nucleus (L.Alvarez, private communication 1973). We have calculated the error assuming that the orbital n is 4 or larger.
- <sup>3</sup>Excluded from the average; the  $\Omega^-$  lifetimes measured by the experiments differ significantly from other measurements.
- <sup>4</sup>The EISENBERG 54 mass was calculated for decay in flight. ALVAREZ 73 has shown that the  $\Omega$  interacted with an Ag nucleus to give  $K^- \Xi \text{Ag}$ .

$\overline{\Omega}^+$  MASS

The fit assumes the  $\Omega^-$  and  $\overline{\Omega}^+$  masses are the same, and averages them together.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1672.45 ± 0.29 OUR FIT</b>				
<b>1672.5 ± 0.7 OUR AVERAGE</b>				
1672 ± 1	72	HARTOUNI 85	SPEC	80–280 GeV $K_L^0 C$
1673.1 ± 1.0	1	FIRESTONE 71B	HBC	12 GeV/c $K^+ d$

$(m_{\Omega^-} - m_{\overline{\Omega}^+}) / m_{\Omega^-}$

A test of  $CPT$  invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>(−1.44 ± 7.98) × 10<sup>−5</sup></b>	CHAN 98	E756	$p$ Be, 800 GeV

$\Omega^-$  MEAN LIFE

Measurements with an error  $> 0.1 \times 10^{-10}$  s have been omitted. The fit assumes the  $\Omega^-$  and  $\overline{\Omega}^+$  mean lives are the same, and averages them together.

VALUE (10 <sup>−10</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.821 ± 0.011 OUR FIT</b>				
<b>0.821 ± 0.011 OUR AVERAGE</b>				
0.817 ± 0.013 ± 0.018	6934	CHAN 98	E756	$p$ Be, 800 GeV
0.811 ± 0.037	1096	LUK 88	SPEC	$p$ Be 400 GeV
0.823 ± 0.013	12k	BOURQUIN 84	SPEC	SPS hyperon beam
• • • We do not use the following data for averages, fits, limits, etc. • • •				

0.822 ± 0.028 2437 BOURQUIN 79B SPEC See BOURQUIN 84

$\overline{\Omega}^+$  MEAN LIFE

The fit assumes the  $\Omega^-$  and  $\overline{\Omega}^+$  mean lives are the same, and averages them together.

VALUE (10 <sup>−10</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.821 ± 0.011 OUR FIT</b>				
<b>0.823 ± 0.031 ± 0.022</b>	1801	CHAN 98	E756	$p$ Be, 800 GeV

$(\tau_{\Omega^-} - \tau_{\overline{\Omega}^+}) / \tau_{\Omega^-}$

A test of  $CPT$  invariance. Our calculation, from the averages in the preceding two data blocks.

VALUE	DOCUMENT ID
<b>0.00 ± 0.05 OUR ESTIMATE</b>	

$\Omega^-$  MAGNETIC MOMENT

VALUE ( $\mu_N$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>−2.02 ± 0.05 OUR AVERAGE</b>				
−2.024 ± 0.056	235k	WALLACE 95	SPEC	$\Omega^-$ 300–550 GeV
−1.94 ± 0.17 ± 0.14	25k	DIEHL 91	SPEC	Spin-transfer production

$\Omega^-$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 \Lambda K^-$	(67.8 ± 0.7) %	
$\Gamma_2 \Xi^0 \pi^-$	(23.6 ± 0.7) %	
$\Gamma_3 \Xi^- \pi^0$	( 8.6 ± 0.4) %	
$\Gamma_4 \Xi^- \pi^+ \pi^-$	( 3.7 ± 0.7 ) × 10 <sup>−4</sup>	
$\Gamma_5 \Xi (1530)^0 \pi^-$	< 7 × 10 <sup>−5</sup>	90%
$\Gamma_6 \Xi^0 e^- \overline{\nu}_e$	( 5.6 ± 2.8 ) × 10 <sup>−3</sup>	
$\Gamma_7 \Xi^- \gamma$	< 4.6 × 10 <sup>−4</sup>	90%

$\Delta S = 2$  forbidden ( $S_2$ ) modes

$\Gamma_8 \Lambda \pi^-$	$S_2$	< 2.9 × 10 <sup>−6</sup>	90%
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$\Omega^-$  BRANCHING RATIOS

The BOURQUIN 84 values (which include results of BOURQUIN 79B, a separate experiment) are much more accurate than any other results, and so the other results have been omitted.

$\Gamma(AK^-)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.678 ± 0.007</b>	14k	BOURQUIN 84	SPEC	SPS hyperon beam
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.686 ± 0.013	1920	BOURQUIN 79B	SPEC	See BOURQUIN 84

$\Gamma(\Xi^0 \pi^-)/\Gamma_{\text{total}}$				$\Gamma_2/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.236 ± 0.007</b>	1947	BOURQUIN 84	SPEC	SPS hyperon beam
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.234 ± 0.013	317	BOURQUIN 79B	SPEC	See BOURQUIN 84

$\Gamma(\Xi^- \pi^0)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.086 ± 0.004</b>	759	BOURQUIN 84	SPEC	SPS hyperon beam
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.080 ± 0.008	145	BOURQUIN 79B	SPEC	See BOURQUIN 84

$\Gamma(\Xi^-\pi^+\pi^-)/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.74^{+0.67}_{-0.56}</math></b>	100	<sup>5</sup> KAMAEV	10	HYCP $p$ Cu, 800 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$4.3^{+3.4}_{-1.3}$	4	BOURQUIN	84	SPEC SPS hyperon beam

- <sup>5</sup> This KAMAEV 10 value uses 76  $\Omega^- \rightarrow \Xi^- \pi^+ \pi^-$  and 24  $\overline{\Omega}^+ \rightarrow \Xi^+ \pi^- \pi^+$  decays. The  $\Omega^-$  and  $\overline{\Omega}^+$  branching fractions measurements are statistically equal. The errors given combine statistical and systematic contributions. The  $CP$  branching-fraction asymmetry,  $(\Omega^- - \overline{\Omega}^+)/\text{sum}$ , is  $+0.12 \pm 0.20$ .



See key on page 601

Baryon Particle Listings

$\Omega^-$ ,  $\Omega(2250)^-$ ,  $\Omega(2380)^-$

$\Gamma(\Xi(1530)^0\pi^-)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.7</b>	90		KAMAEV	10	HYCP $p$ Cu, 800 GeV	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$6.4^{+5.1}_{-2.0}$	4	6	BOURQUIN	84	SPEC	SPS hyperon beam
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<sup>6</sup> The same 4 events as in the previous mode, with the isospin factor to take into account  $\Xi(1530)^0 \rightarrow \Xi^0\pi^0$  decays included. BOURQUIN 84 adopted a theoretical assumption that  $\Xi(1530)^0\pi^-$  would dominate  $\Xi^-\pi^+\pi^-$  decay.

$\Gamma(\Xi^0 e^- \bar{\nu}_e)/\Gamma_{\text{total}}$					$\Gamma_6/\Gamma$	
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b><math>5.6 \pm 2.8</math></b>	14		BOURQUIN	84	SPEC	SPS hyperon beam

• • • We do not use the following data for averages, fits, limits, etc. • • •

~ 10	3		BOURQUIN	79B	SPEC	See BOURQUIN 84
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$\Gamma(\Xi^-\gamma)/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$	
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 4.6</b>	90	0	ALBUQUERQ...94	E761	$\Omega^-$ 375 GeV	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<22	90	9	BOURQUIN	84	SPEC	SPS hyperon beam
<31	90	0	BOURQUIN	79B	SPEC	See BOURQUIN 84

$\Gamma(\Lambda\pi^-)/\Gamma_{\text{total}}$					$\Gamma_8/\Gamma$
$\Delta S=2$ . Forbidden in first-order weak interaction.					
VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt; 2.9</b>	90	WHITE	05	HYCP	$p$ Cu, 800 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 190	90	BOURQUIN	84	SPEC	SPS hyperon beam
<1300	90	BOURQUIN	79B	SPEC	See BOURQUIN 84

$\Omega^-$  DECAY PARAMETERS

$\alpha$  FOR  $\Omega^- \rightarrow \Lambda K^-$

Some early results have been omitted.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.0180 \pm 0.0024</math> OUR AVERAGE</b>				
$+0.0207 \pm 0.0051 \pm 0.0081$	960k	<sup>7</sup> CHEN	05	HYCP $p$ Cu, 800 GeV
$+0.0178 \pm 0.0019 \pm 0.0016$	4.5M	<sup>7</sup> LU	05A	HYCP $p$ Cu, 800 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.028 \pm 0.047$	6953	CHAN	98	E756 $p$ Be, 800 GeV
$-0.034 \pm 0.079$	1743	LUK	88	SPEC $p$ Be 400 GeV
$-0.025 \pm 0.028$	12k	BOURQUIN	84	SPEC SPS hyperon beam

<sup>7</sup> The results of CHEN 05 and LU 05A are from different experimental runs.

$\bar{\alpha}$  FOR  $\bar{\Omega}^+ \rightarrow \bar{\Lambda} K^+$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>-0.0181 \pm 0.0028 \pm 0.0026</math></b>	1.89M	LU	06	HYCP $p$ Cu, 800 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$+0.017 \pm 0.077$	1823	CHAN	98	E756 $p$ Be, 800 GeV
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$(\alpha + \bar{\alpha})/(\alpha - \bar{\alpha})$  in  $\Omega^- \rightarrow \Lambda K^-, \bar{\Omega}^+ \rightarrow \bar{\Lambda} K^+$

Zero if  $CP$  is conserved.

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.016 \pm 0.092 \pm 0.089</math></b>	<sup>8</sup> LU	06	HYCP $p$ Cu, 800 GeV

<sup>8</sup> This value uses the results of CHEN 05, LU 05A, and LU 06.

$\alpha$  FOR  $\Omega^- \rightarrow \Xi^0\pi^-$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>+0.09 \pm 0.14</math></b>	1630	BOURQUIN	84	SPEC SPS hyperon beam

$\alpha$  FOR  $\Omega^- \rightarrow \Xi^-\pi^0$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>+0.05 \pm 0.21</math></b>	614	BOURQUIN	84	SPEC SPS hyperon beam

$\Omega^-$  REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

KAMAEV	10	PL B693 236	O. Kamaev <i>et al.</i>	(FNAL HyperCP Collab.)
AUBERT, BE	06	PRL 97 112001	B. Aubert <i>et al.</i>	(BABAR Collab.)
LU	06	PRL 96 242001	L.C. Lu <i>et al.</i>	(FNAL HyperCP Collab.)
CHEN	05	PR D71 051102	Y.C. Chen <i>et al.</i>	(FNAL HyperCP Collab.)
LU	05A	PL B617 11	L.C. Lu <i>et al.</i>	(FNAL HyperCP Collab.)
WHITE	05	PRL 94 101804	C.G. White <i>et al.</i>	(FNAL HyperCP Collab.)
CHAN	98	PR D58 072002	A.W. Chan <i>et al.</i>	(FNAL E756 Collab.)
WALLACE	95	PRL 74 3732	N.B. Wallace <i>et al.</i>	(MINN, ARIZ, MICH+)
ALBUQUERQ...	94	PR D50 R18	I.F. Albuquerque <i>et al.</i>	(FNAL E761 Collab.)
DIEHL	91	PRL 67 804	H.T. Diehl <i>et al.</i>	(RUTG, FNAL, MICH+)
LUK	88	PR D38 19	K.B. Luk <i>et al.</i>	(RUTG, WISC, MICH, MINN)
HARTOUNI	85	PRL 54 628	E.P. Hartouni <i>et al.</i>	(COLU, ILL, FNAL)
BOURQUIN	84	NP B241 1	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
Also		PL B7B 297	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
BOURQUIN	79B	PL B8B 192	M.H. Bourquin <i>et al.</i>	(BRIS, GEVA, HEIDP+)
BAUBILLIER	78	PL 78B 342	M. Baubillier <i>et al.</i>	(BIRM, CERN, GLAS+)

DEUTSCH...	78	PL 73B 96	M. Deuschmann <i>et al.</i>	(AACH3, BERL, CERN+)
HEMINGWAY	78	NP B142 205	R.J. Hemingway <i>et al.</i>	(CERN, ZEEM, NUM+)
DIBIAN CA	75	NP B98 137	F.A. Dibianca, R.J. Endorf	(CMU)
ALVAREZ	73	PR D8 702	L.W. Alvarez	(LBL)
DEUTSCH...	73	NP B61 102	M. Deuschmann <i>et al.</i>	(ABCLV Collab.)
FIRESTONE	71B	PRL 26 410	I. Firestone <i>et al.</i>	(LRL)
SPETH	69	PL 29B 252	R. Speth <i>et al.</i>	(AACH, BERL, CERN, LOIC+)
PALMER	68	PL 26B 323	R.B. Palmer <i>et al.</i>	(BNL, SYRA)
SCHULTZ	68	PR 168 1509	P.F. Schultz <i>et al.</i>	(ILL, ANL, NWES+)
SCOTTER	68	PL 26B 474	D. Scotter <i>et al.</i>	(BIRM, GLAS, LOIC+)
ABRAMS	64	PRL 13 670	G.S. Abrams <i>et al.</i>	(UMD, NRL)
BARNES	64	PRL 12 204	V.E. Barnes <i>et al.</i>	(BNL)
FRY	55	PR 97 1189	W.F. Fry, J. Schneps, M.S. Swami	(WISC)
FRY	55B	NC 2 346	W.F. Fry, J. Schneps, M.S. Swami	(WISC)
EISENBERG	54	PR 96 541	Y. Eisenberg	(CORN)

$\Omega(2250)^-$	$I(J^P) = 0(?^?)$	Status: ***
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$\Omega(2250)^-$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2252 \pm 9</math> OUR AVERAGE</b>				
$2253 \pm 13$	44	ASTON	87B	LASS $K^- p$ 11 GeV/c
$2251 \pm 9 \pm 8$	78	BIAGI	86B	SPEC SPS $\Xi^-$ beam

$\Omega(2250)^-$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>55 \pm 18</math> OUR AVERAGE</b>				
$81 \pm 38$	44	ASTON	87B	LASS $K^- p$ 11 GeV/c
$48 \pm 20$	78	BIAGI	86B	SPEC SPS $\Xi^-$ beam

$\Omega(2250)^-$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Xi^- \pi^+ K^-$	seen
$\Gamma_2 \Xi(1530)^0 K^-$	seen

$\Omega(2250)^-$  BRANCHING RATIOS

$\Gamma(\Xi(1530)^0 K^-)/\Gamma(\Xi^- \pi^+ K^-)$					$\Gamma_2/\Gamma_1$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
~ 1.0	44	ASTON	87B	LASS $K^- p$ 11 GeV/c		
$0.70 \pm 0.20$	49	BIAGI	86B	SPEC $\Xi^-$ Be 116 GeV/c		

$\Omega(2250)^-$  REFERENCES

ASTON	87B	PL B194 579	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
BIAGI	86B	ZPHY C31 33	S.F. Biagi <i>et al.</i>	(LOQM, GEVA, RAL+)

$\Omega(2380)^-$	Status: **
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OMITTED FROM SUMMARY TABLE

$\Omega(2380)^-$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>\approx 2380</math> OUR ESTIMATE</b>				
$2384 \pm 9 \pm 8$	45	BIAGI	86B	SPEC SPS $\Xi^-$ beam

$\Omega(2380)^-$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
$26 \pm 23$	45	BIAGI	86B	SPEC SPS $\Xi^-$ beam

$\Omega(2380)^-$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Xi^- \pi^+ K^-$	seen
$\Gamma_2 \Xi(1530)^0 K^-$	
$\Gamma_3 \Xi^- \bar{K}^*(892)^0$	

$\Omega(2380)^-$  BRANCHING RATIOS

$\Gamma(\Xi(1530)^0 K^-)/\Gamma(\Xi^- \pi^+ K^-)$					$\Gamma_2/\Gamma_1$	
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<0.44	90	9	BIAGI	86B	SPEC $\Xi^-$ Be 116 GeV/c	

Baryon Particle Listings

$\Omega(2380)^-, \Omega(2470)^-$

$\Gamma(\Xi^-\bar{K}^*(892)^0)/\Gamma(\Xi^-\pi^+K^-)$					$\Gamma_3/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
0.5 ± 0.3	21	BIAGI	86B	SPEC	$\Xi^-$ Be 116 GeV/c

$\Omega(2380)^-$  REFERENCES

BIAGI	86B	ZPHY C31 33	S.F. Biagi <i>et al.</i>	(LOQM, GEVA, RAL+)
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<div><math>\Omega(2470)^-</math></div>	Status: **
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OMITTED FROM SUMMARY TABLE

A peak in the  $\Omega^-\pi^+\pi^-$  mass spectrum with a signal significance claimed to be at least 5.5 standard deviations. There is no reason to seriously doubt the existence of this state, but unless the evidence is overwhelming we usually wait for confirmation from a second experiment before elevating peaks to the Summary Table.

$\Omega(2470)^-$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
2474 ± 12	59	ASTON	88G	LASS $K^-p$ 11 GeV/c

$\Omega(2470)^-$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
72 ± 33	59	ASTON	88G	LASS $K^-p$ 11 GeV/c

$\Omega(2470)^-$  DECAY MODES

Mode
$\Gamma_1 \quad \Omega^-\pi^+\pi^-$

$\Omega(2470)^-$  REFERENCES

ASTON	88G	PL B215 799	D. Aston <i>et al.</i>	(SLAC, NAGO, CINC, INUS)
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## CHARMED BARYONS ( $C = +1$ )

$$\begin{aligned}\Lambda_c^+ &= udc, & \Sigma_c^{++} &= uuc, & \Sigma_c^+ &= udc, & \Sigma_c^0 &= ddc, \\ \Xi_c^+ &= usc, & \Xi_c^0 &= dsc, & \Omega_c^0 &= ssc\end{aligned}$$

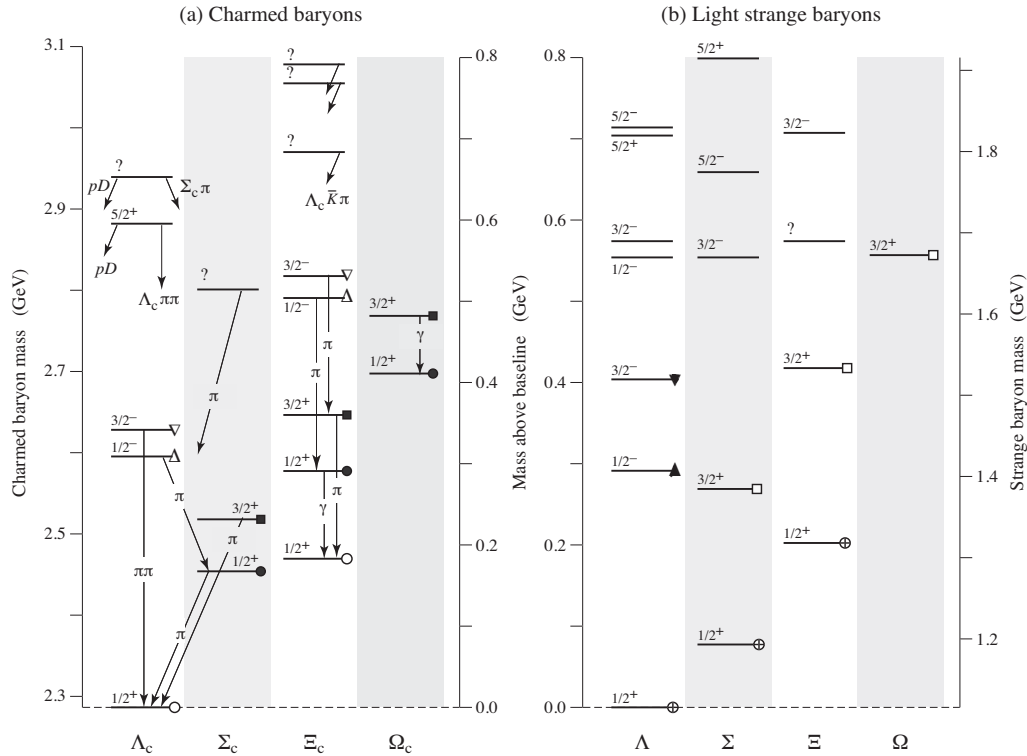
### CHARMED BARYONS

Revised March 2012 by C.G. Wohl (LBNL).

(Note added November 2015.) Since the 2014 *Review*, there have been three papers that improved the values of the masses and/or widths of the  $\Sigma(2455)$ ,  $\Sigma(2520)$ ,  $\Xi_c^+$ ,  $\Xi_c^0$ ,  $\Xi_c(2645)$ ,  $\Xi_c(2970)$ ,  $\Xi_c(3055)$ , and  $\Xi_c(3080)$ . See the Listings for those particles.

There are 18 known charmed baryons, and four other candidates not well enough established to be promoted to the Summary Tables.\* Fig. 1(a) shows the mass spectrum, and for comparison Fig. 1(b) shows the spectrum of the lightest strange baryons. The  $\Lambda_c$  and  $\Sigma_c$  spectra ought to look much like the  $\Lambda$  and  $\Sigma$  spectra, since a  $\Lambda_c$  or a  $\Sigma_c$  differs from a  $\Lambda$  or a  $\Sigma$  only by the replacement of the  $s$  quark with a  $c$  quark. However, a  $\Xi$  or an  $\Omega$  has more than one  $s$  quark, only *one* of which is changed to a  $c$  quark to make a  $\Xi_c$  or an  $\Omega_c$ . Thus the  $\Xi_c$  and  $\Omega_c$  spectra ought to be richer than the  $\Xi$  and  $\Omega$  spectra.\*\*

Before discussing the observed spectra, we review the theory of SU(4) multiplets, which tells what charmed baryons to expect; this is essential, because few of the spin-parity values given in Fig. 1(a) have been measured.



**Fig. 1.** (a) The known charmed baryons, and (b) the lightest “4-star” strange baryons. Note that there are two  $J^P = 1/2^+$   $\Xi_c$  states, and that the lightest  $\Omega_c$  does not have  $J = 3/2$ . The  $J^P = 1/2^+$  states, all tabbed with a circle, belong to the SU(4) multiplet that includes the nucleon; states with a circle and the same fill belong to the same SU(3) multiplet within that SU(4) multiplet. Similar remarks apply to the other states: same shape of tab, same SU(4) multiplet; same fill of that shape, same SU(3) multiplet. The  $J^P = 1/2^-$  and  $3/2^-$  states tabbed with triangles complete two SU(4) 4 multiplets.

But far and away the most important advance for  $c$ -baryons this edition is the first ever (!) model-independent measurement of a  $\Lambda_c^+$  branching fraction. The anchor for all the other  $\Lambda_c^+$  fractions is the  $pK^-\pi^+$  fraction. Our value for many editions,  $(5.0 \pm 1.3)\%$ , was cobbled together from two ancient model-dependent measurements. Now, thanks to the Belle experiment, it is  $(6.84 \pm 0.24_{\pm 0.27})\%$ . This fraction is 37% larger than the old fraction, and its error is only 27% as large as the old error.

Rather, they have been assigned in accord with expectations of the theory. However, they are all very likely as shown (see below).

**SU(4) multiplets**—Baryons made from  $u$ ,  $d$ ,  $s$ , and  $c$  quarks belong to SU(4) multiplets. The multiplet numerology, analogous to  $3 \times 3 \times 3 = 10 + 8_1 + 8_2 + 1$  for the subset of baryons made from just  $u$ ,  $d$ , and  $s$  quarks, is  $4 \times 4 \times 4 = 20 + 20'_1 + 20'_2 + \bar{4}$ . Figure 2(a) shows the 20-plet whose bottom level is an SU(3) decuplet, such as the decuplet that includes the  $\Delta(1232)$ .

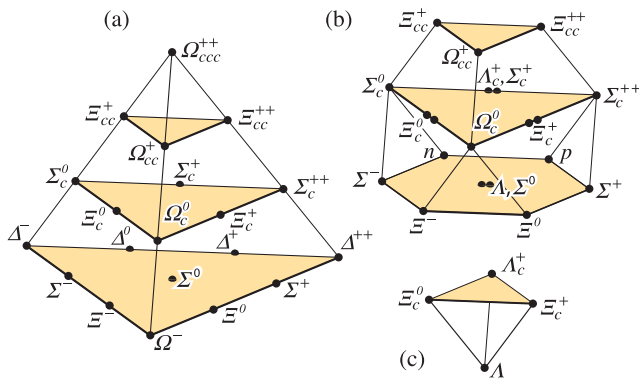
Figure 2(b) shows the 20'-plet whose bottom level is an SU(3) octet, such as the octet that includes the nucleon.

## Baryon Particle Listings

### Charmed Baryons

Figure 2(c) shows the  $\bar{4}$  multiplet, an inverted tetrahedron. One level up from the bottom level of each multiplet are the baryons with one  $c$  quark. All the baryons in a given multiplet have the same spin and parity. Each  $N$  or  $\Delta$  or  $SU(3)$ -singlet- $\Lambda$  resonance calls for another  $20'$ - or  $20$ - or  $4$ -plet, respectively.

The flavor symmetries shown in Fig. 2 are of course badly broken, but the figure is the simplest way to see what charmed baryons should exist. For example, from Fig. 2(b), we expect to find, in the same  $J^P = 1/2^+$   $20'$ -plet as the nucleon, a  $\Lambda_c$ , a  $\Sigma_c$ , two  $\Xi_c$ 's, and an  $\Omega_c$ . Note that this  $\Omega_c$  has  $J^P = 1/2^+$  and is not in the same  $SU(4)$  multiplet as the famous  $J^P = 3/2^+$   $\Omega^-$ .

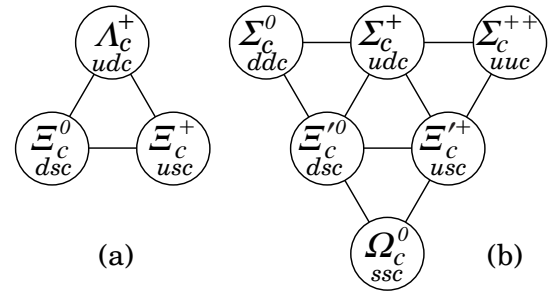


**Figure 2:**  $SU(4)$  multiplets of baryons made of  $u$ ,  $d$ ,  $s$ , and  $c$  quarks. (a) The  $20$ -plet with an  $SU(3)$  decuplet on the lowest level. (b) The  $20'$ -plet with an  $SU(3)$  octet on the lowest level. (c) The  $4$ -plet. Note that here and in Fig. 3, but not in Fig. 1, each charge state is shown separately.

Figure 3 shows in more detail the middle level of the  $20'$ -plet of Fig. 2(b); it splits apart into two  $SU(3)$  multiplets, a  $\bar{3}$  and a  $6$ . The states of the  $\bar{3}$  are antisymmetric under the interchange of the two light quarks (the  $u$ ,  $d$ , and  $s$  quarks), whereas the states of the  $6$  are symmetric under this interchange. We use a prime to distinguish the  $\Xi_c$  in the  $6$  from the one in the  $\bar{3}$ .

**The observed spectra**—(1) The parity of the lightest  $\Lambda_c$  is defined to be positive (as are the parities of the  $p$ ,  $n$ , and  $\Lambda$ ); the limited evidence about its spin is consistent with  $J = 1/2$ . However, few of the  $J^P$  quantum numbers given in Fig. 1(a) have been measured. Models using spin-spin and spin-orbit interactions between the quarks, with parameters determined using a few of the masses as input, lead to the  $J^P$  assignments shown.<sup>†</sup> There are no surprises: the  $J^P = 1/2^+$  states come first, then the  $J^P = 3/2^+$  states...

(2) There is, however, evidence that many of the  $J^P$  assignments in Fig. 1(a) must be correct. As is well known, the successive mass differences between the  $J^P = 3/2^+$  particles, the  $\Delta(1232)^-$ ,  $\Sigma(1385)^-$ ,  $\Xi(1535)^-$ , and  $\Omega^-$ , which lie along the lower left edge of the  $20$ -plet in Fig. 2(a), should according



**Figure 3:** The  $SU(3)$  multiplets on the second level of the  $SU(4)$  multiplet of Fig. 2(b). The  $\Lambda_c$  and  $\Xi_c$  tabbed with open circles in Fig. 1(a) complete a  $J^P = 1/2^+$   $SU(3)$   $\bar{3}$ -plet, as in (a) here. The  $\Sigma_c$ ,  $\Xi_c$ , and  $\Omega_c$  tabbed with closed circles in Fig. 1(a) complete a  $J^P = 1/2^+$   $SU(3)$   $6$ -plet, as in (b) here. Together the nine particles complete the charm = +1 level of a  $J^P = 1/2^+$   $SU(4)$   $20'$ -plet, as in Fig. 2(b).

to  $SU(3)$  be about equal; and indeed experimentally they nearly are. In the same way, the mass differences between the  $J^P = 1/2^+$   $\Sigma_c(2455)^0$ ,  $\Xi_c^0$ , and  $\Omega_c^0$ ,<sup>‡</sup> the particles along the left edge of Fig. 3(b), should be about equal—assuming, of course, that they *do* all have the same  $J^P$ . The measured differences are  $125.0 \pm 2.9$  MeV and  $117.3 \pm 3.4$  MeV—not perfect, but close. Similarly, the mass differences between the presumed  $J^P = 3/2^+$   $\Sigma_c(2520)^0$ ,  $\Xi_c(2645)^0$ , and  $\Omega_c(2770)^0$  are  $127.1 \pm 0.8$  MeV and  $120.0 \pm 2.1$  MeV. In Fig. 1(a), these two sets of charm particles are tabbed with solid circles and solid squares.

(3) Other evidence comes from the decay of the  $\Lambda_c(2593)$ . The only allowed strong decay is  $\Lambda_c(2593)^+ \rightarrow \Lambda_c^+ \pi \pi$ , and this appears to be dominated by the submode  $\Sigma_c(2455)\pi$ , despite little available phase space for the latter (the “ $Q$ ” is about 2 MeV, the c.m. decay momentum about 20 MeV/c). Thus the decay is almost certainly  $s$ -wave, which, assuming that the  $\Sigma_c(2455)$  does indeed have  $J^P = 1/2^+$ , makes  $J^P = 1/2^-$  for the  $\Lambda_c(2593)$ .

#### Footnotes:

\* The unpromoted states are a  $\Lambda_c(2765)^+$ , a  $\Xi_c(2930)$ , and a  $\Xi_c(3123)$ . There is also very weak evidence for a baryon with *two*  $c$  quarks, a  $\Xi_{cc}^+$  at 3519 MeV. See the Particle Listings.

\*\* For example, there are three  $\Omega_c^0$  states (properly symmetrized states of  $ssc$ ,  $scs$ , and  $css$ ) corresponding to each  $\Omega^-$  ( $sss$ ) state.

† This is not the place to discuss the details of the models, nor to attempt a guide to the literature. See the discovery papers of the various charmed baryons for references to the models that lead to the quantum-number assignments.

‡ A reminder about the Particle Data Group naming scheme: A particle has its mass as part of its name if and only if it decays strongly. Thus  $\Sigma(1385)$  and  $\Sigma_c(2455)$  but  $\Omega^-$  and  $\Xi_c'$ .

See key on page 601

## Baryon Particle Listings

 $\Lambda_c^+$ 

$$I(J^P) = 0(\frac{1}{2}^+) \text{ Status: } ****$$

The parity of the  $\Lambda_c^+$  is defined to be positive (as are the parities of the proton, neutron, and  $\Lambda$ ). The quark content is  $udc$ . Results of an analysis of  $pK^-\pi^+$  decays (JEZABEK 92) are consistent with  $J = 1/2$ . Nobody doubts that the spin is indeed  $1/2$ .

We have omitted some results that have been superseded by later experiments. The omitted results may be found in earlier editions.

 $\Lambda_c^+$  MASS

Our value in 2004,  $2284.9 \pm 0.6$  MeV, was the average of the measurements now filed below as "not used." The BABAR measurement is so much better that we use it alone. Note that it is about 2.6 (old) standard deviations above the 2004 value.

The fit also includes  $\Sigma_c^-\Lambda_c^+$  and  $\Lambda_c^{*+}\Lambda_c^+$  mass-difference measurements, but this doesn't affect the  $\Lambda_c^+$  mass. The new (in 2006)  $\Lambda_c^+$  mass simply pushes all those other masses higher.

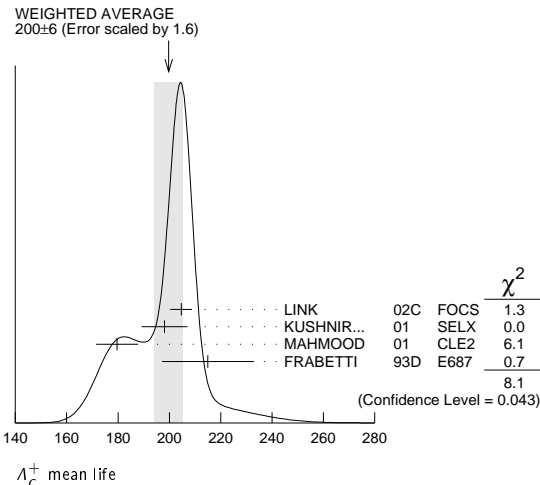
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2286.46 ± 0.14 OUR FIT</b>				
<b>2286.46 ± 0.14</b>	4891	<sup>1</sup> AUBERT,B	05s BABR	$\Lambda_c^0 K^+$ and $\Sigma^0 K_S^0 K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2284.7 ± 0.6 ± 0.7	1134	AVERY	91 CLEO	Six modes
2281.7 ± 2.7 ± 2.6	29	ALVAREZ	90b NA14	$pK^-\pi^+$
2285.8 ± 0.6 ± 1.2	101	BARLAG	89 NA32	$pK^-\pi^+$
2284.7 ± 2.3 ± 0.5	5	AGUILAR...	88b LEBC	$pK^-\pi^+$
2283.1 ± 1.7 ± 2.0	628	ALBRECHT	88c ARG	$pK^-\pi^+$ , $p\bar{K}^0, \Lambda 3\pi$
2286.2 ± 1.7 ± 0.7	97	ANJOS	88b E691	$pK^-\pi^+$
2281 ± 3	2	JONES	87 HBC	$pK^-\pi^+$
2283 ± 3	3	BOSETTI	82 HBC	$pK^-\pi^+$
2290 ± 3	1	CALICCHIO	80 HYBR	$pK^-\pi^+$

<sup>1</sup> AUBERT,B 05s uses low-Q  $\Lambda_c^0 K^+$  and  $\Sigma^0 K_S^0 K^+$  decays to minimize systematic errors. The error above includes systematic as well as statistical errors. Many cross checks and adjustments to properties of the BABAR detector, as well as the large number of clean events, make this by far the best measurement of the  $\Lambda_c^+$  mass.

 $\Lambda_c^+$  MEAN LIFE

Measurements with an error  $\geq 100 \times 10^{-15}$  s or with fewer than 20 events have been omitted from the Listings.

VALUE ( $10^{-15}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>200 ± 6 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
204.6 ± 3.4 ± 2.5	8034	LINK	02c FOCS	$pK^-\pi^+$
198.1 ± 7.0 ± 5.6	1630	KUSHNIR...	01 SELX	$\Lambda_c^+ \rightarrow pK^-\pi^+$
179.6 ± 6.9 ± 4.4	4749	MAHMOOD	01 CLE2	$e^+e^- \approx \tau(4S)$
215 ± 16 ± 8	1340	FRABETTI	93D E687	$\gamma\text{Be}, \Lambda_c^+ \rightarrow pK^-\pi^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
180 ± 30 ± 30	29	ALVAREZ	90 NA14	$\gamma, \Lambda_c^+ \rightarrow pK^-\pi^+$
200 ± 30 ± 30	90	FRABETTI	90 E687	$\gamma\text{Be}, \Lambda_c^+ \rightarrow pK^-\pi^+$
196 $^{+23}_{-20}$	101	BARLAG	89 NA32	$pK^-\pi^+$ + c.c.
220 ± 30 ± 20	97	ANJOS	88b E691	$pK^-\pi^+$ + c.c.

 $\Lambda_c^+$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Hadronic modes with a <math>p</math>: <math>S = -1</math> final states</b>		
$\Gamma_1$ $pK_S^0$	( 1.58 ± 0.08 ) %	S=1.2
$\Gamma_2$ $pK^-\pi^+$	( 6.35 ± 0.33 ) %	S=1.4
$\Gamma_3$ $p\bar{K}^*(892)^0$	[a] ( 1.98 ± 0.28 ) %	
$\Gamma_4$ $\Delta(1232)^{++}K^-$	( 1.09 ± 0.25 ) %	
$\Gamma_5$ $\Lambda(1520)\pi^+$	[a] ( 2.2 ± 0.5 ) %	
$\Gamma_6$ $pK^-\pi^+$ nonresonant	( 3.5 ± 0.4 ) %	
$\Gamma_7$ $pK_S^0\pi^0$	( 1.99 ± 0.13 ) %	S=1.1
$\Gamma_8$ $p\bar{K}^0\eta$	( 1.6 ± 0.4 ) %	
$\Gamma_9$ $pK_S^0\pi^+\pi^-$	( 1.66 ± 0.12 ) %	S=1.1
$\Gamma_{10}$ $pK^-\pi^+\pi^0$	( 4.9 ± 0.4 ) %	S=1.3
$\Gamma_{11}$ $pK^*(892)^-\pi^+$	[a] ( 1.5 ± 0.5 ) %	
$\Gamma_{12}$ $p(K^-\pi^+)_{\text{nonresonant}}\pi^0$	( 4.6 ± 0.9 ) %	
$\Gamma_{13}$ $\Delta(1232)\bar{K}^*(892)$	seen	
$\Gamma_{14}$ $pK^-\pi^+\pi^-\pi^0$	( 1.4 ± 1.0 ) $\times 10^{-3}$	
$\Gamma_{15}$ $pK^-\pi^+2\pi^0$	( 1.0 ± 0.5 ) %	
$\Gamma_{16}$ $pK^-\pi^+3\pi^0$		
<b>Hadronic modes with a <math>p</math>: <math>S = 0</math> final states</b>		
$\Gamma_{17}$ $p\pi^+\pi^-$	( 4.4 ± 2.3 ) $\times 10^{-3}$	
$\Gamma_{18}$ $p\bar{f}_0(980)$	[a] ( 3.5 ± 2.3 ) $\times 10^{-3}$	
$\Gamma_{19}$ $2\pi^+2\pi^-$	( 2.3 ± 1.5 ) $\times 10^{-3}$	
$\Gamma_{20}$ $pK^+K^-$	( 10 ± 4 ) $\times 10^{-4}$	
$\Gamma_{21}$ $p\phi$	[a] ( 1.04 ± 0.21 ) $\times 10^{-3}$	
$\Gamma_{22}$ $pK^+K^-$ non- $\phi$	( 4.4 ± 1.8 ) $\times 10^{-4}$	
<b>Hadronic modes with a hyperon: <math>S = -1</math> final states</b>		
$\Gamma_{23}$ $\Lambda\pi^+$	( 1.30 ± 0.07 ) %	S=1.2
$\Gamma_{24}$ $\Lambda\pi^+\pi^0$	( 7.1 ± 0.4 ) %	S=1.2
$\Gamma_{25}$ $\Lambda\rho^+$	< 6 %	CL=95%
$\Gamma_{26}$ $\Lambda\pi^-2\pi^+$	( 3.7 ± 0.4 ) %	S=1.9
$\Gamma_{27}$ $\Sigma(1385)^+\pi^+\pi^-, \Sigma^{*+} \rightarrow \Lambda\pi^+$	( 1.0 ± 0.5 ) %	
$\Gamma_{28}$ $\Sigma(1385)^-2\pi^+, \Sigma^{*-} \rightarrow \Lambda\pi^-$	( 7.8 ± 1.6 ) $\times 10^{-3}$	
$\Gamma_{29}$ $\Lambda\pi^+\rho^0$	( 1.5 ± 0.6 ) %	
$\Gamma_{30}$ $\Sigma(1385)^+\rho^0, \Sigma^{*+} \rightarrow \Lambda\pi^+$	( 5 ± 4 ) $\times 10^{-3}$	
$\Gamma_{31}$ $\Lambda\pi^-2\pi^+$ nonresonant	< 1.1 %	CL=90%
$\Gamma_{32}$ $\Lambda\pi^-\pi^02\pi^+$ total	( 2.3 ± 0.8 ) %	
$\Gamma_{33}$ $\Lambda\pi^+\eta$	[a] ( 2.3 ± 0.5 ) %	
$\Gamma_{34}$ $\Sigma(1385)^+\eta$	[a] ( 1.08 ± 0.32 ) %	
$\Gamma_{35}$ $\Lambda\pi^+\omega$	[a] ( 1.5 ± 0.5 ) %	
$\Gamma_{36}$ $\Lambda\pi^-\pi^02\pi^+, \text{ no } \eta \text{ or } \omega$	< 8 $\times 10^{-3}$	CL=90%
$\Gamma_{37}$ $\Lambda K^+\bar{K}^0$	( 5.7 ± 1.1 ) $\times 10^{-3}$	S=2.0
$\Gamma_{38}$ $\Xi(1690)^0K^+, \Xi^{*0} \rightarrow \Lambda\bar{K}^0$	( 1.6 ± 0.5 ) $\times 10^{-3}$	
$\Gamma_{39}$ $\Sigma^0\pi^+$	( 1.29 ± 0.07 ) %	S=1.1
$\Gamma_{40}$ $\Sigma^+\pi^0$	( 1.24 ± 0.10 ) %	
$\Gamma_{41}$ $\Sigma^+\eta$	( 7.0 ± 2.3 ) $\times 10^{-3}$	
$\Gamma_{42}$ $\Sigma^+\pi^+\pi^-$	( 4.57 ± 0.29 ) %	S=1.2
$\Gamma_{43}$ $\Sigma^+\rho^0$	< 1.7 %	CL=95%
$\Gamma_{44}$ $\Sigma^-2\pi^+$	( 2.1 ± 0.4 ) %	
$\Gamma_{45}$ $\Sigma^0\pi^+\pi^0$	( 2.3 ± 0.9 ) %	
$\Gamma_{46}$ $\Sigma^0\pi^-2\pi^+$	( 1.13 ± 0.29 ) %	
$\Gamma_{47}$ $\Sigma^+\pi^+\pi^-\pi^0$	—	
$\Gamma_{48}$ $\Sigma^+\omega$	[a] ( 1.74 ± 0.21 ) %	
$\Gamma_{49}$ $\Sigma^+K^+K^-$	( 3.6 ± 0.4 ) $\times 10^{-3}$	
$\Gamma_{50}$ $\Sigma^+\phi$	[a] ( 4.0 ± 0.6 ) $\times 10^{-3}$	S=1.1
$\Gamma_{51}$ $\Xi(1690)^0K^+, \Xi^{*0} \rightarrow \Sigma^+K^-$	( 1.03 ± 0.26 ) $\times 10^{-3}$	
$\Gamma_{52}$ $\Sigma^+K^+K^-$ nonresonant	< 8 $\times 10^{-4}$	CL=90%
$\Gamma_{53}$ $\Xi^0K^+$	( 5.0 ± 1.2 ) $\times 10^{-3}$	
$\Gamma_{54}$ $\Xi^-K^+\pi^+$	( 6.2 ± 0.6 ) $\times 10^{-3}$	S=1.1
$\Gamma_{55}$ $\Xi(1530)^0K^+$	[a] ( 3.3 ± 0.9 ) $\times 10^{-3}$	
<b>Hadronic modes with a hyperon: <math>S = 0</math> final states</b>		
$\Gamma_{56}$ $\Lambda K^+$	( 6.1 ± 1.2 ) $\times 10^{-4}$	
$\Gamma_{57}$ $\Lambda K^+\pi^+\pi^-$	< 5 $\times 10^{-4}$	CL=90%
$\Gamma_{58}$ $\Sigma^0K^+$	( 5.2 ± 0.8 ) $\times 10^{-4}$	
$\Gamma_{59}$ $\Sigma^0K^+\pi^+\pi^-$	< 2.6 $\times 10^{-4}$	CL=90%
$\Gamma_{60}$ $\Sigma^+K^+\pi^-$	( 2.1 ± 0.6 ) $\times 10^{-3}$	
$\Gamma_{61}$ $\Sigma^+K^*(892)^0$	[a] ( 3.6 ± 1.0 ) $\times 10^{-3}$	
$\Gamma_{62}$ $\Sigma^-K^+\pi^+$	< 1.2 $\times 10^{-3}$	CL=90%
<b>Doubly Cabibbo-suppressed modes</b>		
$\Gamma_{63}$ $pK^+\pi^-$	< 2.9 $\times 10^{-4}$	CL=90%

# Baryon Particle Listings

$\Lambda_C^+$

Semileptonic modes

$\Gamma_{64}$	$\Lambda e^+ \nu_e$	( 3.6 ± 0.4 ) %
$\Gamma_{65}$	$\Lambda \mu^+ \nu_\mu$	

Inclusive modes

$\Gamma_{66}$	$e^+$ anything	( 4.5 ± 1.7 ) %	
$\Gamma_{67}$	$p e^+$ anything	( 1.8 ± 0.9 ) %	
$\Gamma_{68}$	$\Lambda e^+$ anything		
$\Gamma_{69}$	$p$ anything	(50 ± 16 ) %	
$\Gamma_{70}$	$p$ anything (no $\Lambda$ )	(12 ± 19 ) %	
$\Gamma_{71}$	$p$ hadrons		
$\Gamma_{72}$	$n$ anything	(50 ± 16 ) %	
$\Gamma_{73}$	$n$ anything (no $\Lambda$ )	(29 ± 17 ) %	
$\Gamma_{74}$	$\Lambda$ anything	(35 ± 11 ) %	S=1.4
$\Gamma_{75}$	$\Sigma^\pm$ anything	[b] (10 ± 5 ) %	
$\Gamma_{76}$	3prongs	(24 ± 8 ) %	

$\Delta C = 1$  weak neutral current ( $C1$ ) modes, or  
Lepton Family number ( $LF$ ), or Lepton number ( $L$ ), or  
Baryon number ( $B$ ) violating modes

$\Gamma_{77}$	$p e^+ e^-$	$C1$	< 5.5	$\times 10^{-6}$	CL=90%
$\Gamma_{78}$	$p \mu^+ \mu^-$	$C1$	< 4.4	$\times 10^{-5}$	CL=90%
$\Gamma_{79}$	$p e^+ \mu^-$	$LF$	< 9.9	$\times 10^{-6}$	CL=90%
$\Gamma_{80}$	$p e^- \mu^+$	$LF$	< 1.9	$\times 10^{-5}$	CL=90%
$\Gamma_{81}$	$\bar{p} 2 e^+$	$L, B$	< 2.7	$\times 10^{-6}$	CL=90%
$\Gamma_{82}$	$\bar{p} 2 \mu^+$	$L, B$	< 9.4	$\times 10^{-6}$	CL=90%
$\Gamma_{83}$	$\bar{p} e^+ \mu^+$	$L, B$	< 1.6	$\times 10^{-5}$	CL=90%
$\Gamma_{84}$	$\Sigma^- \mu^+ \mu^+$	$L$	< 7.0	$\times 10^{-4}$	CL=90%

[a] This branching fraction includes all the decay modes of the final-state resonance.

[b] The value is for the sum of the charge states or particle/antiparticle states indicated.

CONSTRAINED FIT INFORMATION

An overall fit to 36 branching ratios uses 57 measurements and one constraint to determine 19 parameters. The overall fit has a  $\chi^2 = 39.9$  for 39 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_2$	50									
$x_7$	43	59								
$x_9$	48	57	37							
$x_{10}$	30	83	48	54						
$x_{23}$	46	72	47	36	53					
$x_{24}$	37	66	44	31	52	69				
$x_{26}$	53	13	13	44	9	15	6			
$x_{37}$	14	24	15	13	18	27	20	4		
$x_{39}$	49	58	40	36	42	74	59	28	20	
$x_{40}$	36	42	32	25	30	35	34	14	11	31
$x_{42}$	42	79	48	52	72	54	52	18	18	45
$x_{44}$	15	29	17	17	24	20	19	4	7	17
$x_{46}$	17	11	8	15	9	9	7	25	3	11
$x_{48}$	19	31	19	24	29	20	19	13	7	17
$x_{49}$	22	41	25	26	37	28	27	9	9	24
$x_{50}$	17	33	20	22	30	23	22	8	8	19
$x_{54}$	26	43	27	22	33	53	38	8	15	40
	$x_1$	$x_2$	$x_7$	$x_9$	$x_{10}$	$x_{23}$	$x_{24}$	$x_{26}$	$x_{37}$	$x_{39}$
$x_{42}$	34									
$x_{44}$	12	26								
$x_{46}$	7	11	3							
$x_{48}$	15	28	9	6						
$x_{49}$	18	49	13	5	14					
$x_{50}$	14	42	11	4	12	20				
$x_{54}$	20	33	12	5	12	17	14			
	$x_{40}$	$x_{42}$	$x_{44}$	$x_{46}$	$x_{48}$	$x_{49}$	$x_{50}$			

$\Lambda_C^+$  BRANCHING RATIOS

A few really obsolete results have been omitted.

Hadronic modes with a  $p$ :  $S = -1$  final states

$\Gamma(p K_S^0) / \Gamma_{\text{total}}$	$\Gamma_1 / \Gamma$
VALUE (%)	EVTS
1.58 ± 0.08 OUR FIT	Error includes scale factor of 1.2.
1.52 ± 0.08 ± 0.03	1243 ABLIKIM 16 BES3 $e^+ e^- \rightarrow \Lambda_C \bar{\Lambda}_C$ , 4.599 GeV

$\Gamma(p K_S^0) / \Gamma(p K^- \pi^+)$	$\Gamma_1 / \Gamma_2$
Measurements given as a $\bar{K}^0$ ratio have been divided by 2 to convert to a $K_S^0$ ratio.	
VALUE	EVTS
0.249 ± 0.013 OUR FIT	Error includes scale factor of 1.5.
0.234 ± 0.020 OUR AVERAGE	
0.23 ± 0.01 ± 0.02	1025 ALAM 98 CLE2 $e^+ e^- \approx \gamma(4S)$
0.22 ± 0.04 ± 0.03	133 AVERY 91 CLEO $e^+ e^-$ 10.5 GeV
0.28 ± 0.09 ± 0.07	45 ANJOS 90 E691 $\gamma$ Be 70–260 GeV
0.31 ± 0.08 ± 0.02	73 ALBRECHT 88c ARG $e^+ e^-$ 10 GeV

$\Gamma(p K^- \pi^+) / \Gamma_{\text{total}}$	$\Gamma_2 / \Gamma$
VALUE (%)	EVTS
6.35 ± 0.33 OUR FIT	Error includes scale factor of 1.4.
6.3 ± 0.5 OUR AVERAGE	Error includes scale factor of 2.0.
5.84 ± 0.27 ± 0.23	6.3k ABLIKIM 16 BES3 $e^+ e^- \rightarrow \Lambda_C \bar{\Lambda}_C$ , 4.599 GeV
6.84 ± 0.24 ± 0.21 − 0.27	1.4k 1 ZUPANC 14 BELL $e^+ e^- \rightarrow D^{(*)} - \bar{p} \pi^+$ recoil
• • • We do not use the following data for averages, fits, limits, etc. • • •	
5.0 ± 1.3	2 PDG 02 See footnote

1 This ZUPANC 14 value is the FIRST-EVER model-independent measurement of a  $\Lambda_C^+$  branching fraction.

2 See the note by P. Burchat, " $\Lambda_C^+$  Branching Fractions," in any edition of the Review from 2002 through 2014 for how this value was obtained. It is now obsolete.

$\Gamma(p \bar{K}^*(892)^0) / \Gamma(p K^- \pi^+)$	$\Gamma_3 / \Gamma_2$
Unseen decay modes of the $\bar{K}^*(892)^0$ are included.	
VALUE	EVTS
0.31 ± 0.04 OUR AVERAGE	
0.29 ± 0.04 ± 0.03	1 AITALA 00 E791 $\pi^- N$ , 500 GeV
0.35 ± 0.06 − 0.07 ± 0.03	39 BOZEK 93 NA 32 $\pi^-$ Cu 230 GeV
0.42 ± 0.24	12 BASILE 81b CNTR $p p \rightarrow \Lambda_C^+ e^- X$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
0.35 ± 0.11	BARLAG 90b NA 32 See BOZEK 93

1 AITALA 00 makes a coherent 5-dimensional amplitude analysis of  $946 \pm 38 \Lambda_C^+ \rightarrow p K^- \pi^+$  decays.

$\Gamma(\Delta(1232)^{++} K^-) / \Gamma(p K^- \pi^+)$	$\Gamma_4 / \Gamma_2$
VALUE	EVTS
0.17 ± 0.04 OUR AVERAGE	Error includes scale factor of 1.1.
0.18 ± 0.03 ± 0.03	1 AITALA 00 E791 $\pi^- N$ , 500 GeV
0.12 ± 0.04 − 0.05 ± 0.05	14 BOZEK 93 NA 32 $\pi^-$ Cu 230 GeV
0.40 ± 0.17	17 BASILE 81b CNTR $p p \rightarrow \Lambda_C^+ e^- X$

1 AITALA 00 makes a coherent 5-dimensional amplitude analysis of  $946 \pm 38 \Lambda_C^+ \rightarrow p K^- \pi^+$  decays.

$\Gamma(\Lambda(1520) \pi^+) / \Gamma(p K^- \pi^+)$	$\Gamma_5 / \Gamma_2$
Unseen decay modes of the $\Lambda(1520)$ are included.	
VALUE	EVTS
0.35 ± 0.08 OUR AVERAGE	
0.34 ± 0.08 ± 0.05	1 AITALA 00 E791 $\pi^- N$ , 500 GeV
0.40 ± 0.18 − 0.13 ± 0.09	12 BOZEK 93 NA 32 $\pi^-$ Cu 230 GeV

1 AITALA 00 makes a coherent 5-dimensional amplitude analysis of  $946 \pm 38 \Lambda_C^+ \rightarrow p K^- \pi^+$  decays.

$\Gamma(p K^- \pi^+ \text{ nonresonant}) / \Gamma(p K^- \pi^+)$	$\Gamma_6 / \Gamma_2$
VALUE	EVTS
0.55 ± 0.06 OUR AVERAGE	
0.55 ± 0.06 ± 0.04	1 AITALA 00 E791 $\pi^- N$ , 500 GeV
0.56 ± 0.07 − 0.09 ± 0.05	71 BOZEK 93 NA 32 $\pi^-$ Cu 230 GeV

1 AITALA 00 makes a coherent 5-dimensional amplitude analysis of  $946 \pm 38 \Lambda_C^+ \rightarrow p K^- \pi^+$  decays.

$\Gamma(p K_S^0 \pi^0) / \Gamma_{\text{total}}$	$\Gamma_7 / \Gamma$
VALUE (%)	EVTS
1.99 ± 0.13 OUR FIT	Error includes scale factor of 1.1.
1.87 ± 0.13 ± 0.05	558 ABLIKIM 16 BES3 $e^+ e^- \rightarrow \Lambda_C \bar{\Lambda}_C$ , 4.599 GeV

See key on page 601

## Baryon Particle Listings

 $\Lambda_c^+$ 

$\Gamma(pK_S^0\pi^0)/\Gamma(pK^-\pi^+)$   $\Gamma_7/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.313±0.018 OUR FIT</b>				
<b>0.33 ±0.03 ±0.04</b>	774	ALAM	98	CLE2 $e^+e^- \approx \Upsilon(4S)$

$\Gamma(p\bar{K}^0\eta)/\Gamma(pK^-\pi^+)$   $\Gamma_8/\Gamma_2$   
Unseen decay modes of the  $\eta$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.25±0.04±0.04</b>	57	AMMAR	95	CLE2 $e^+e^- \approx \Upsilon(4S)$

$\Gamma(pK_S^0\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_9/\Gamma$   
Measurements given as a  $K_S^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1.66±0.12 OUR FIT</b>				Error includes scale factor of 1.1.
<b>1.53±0.11±0.09</b>	485	ABLIKIM	16	BES3 $e^+e^- \rightarrow \Lambda_c\bar{\Lambda}_c$ , 4.599 GeV

$\Gamma(pK_S^0\pi^+\pi^-)/\Gamma(pK^-\pi^+)$   $\Gamma_9/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.261±0.016 OUR FIT</b>				Error includes scale factor of 1.2.
<b>0.257±0.031 OUR AVERAGE</b>				
0.26 ±0.02 ±0.03	985	ALAM	98	CLE2 $e^+e^- \approx \Upsilon(4S)$
0.22 ±0.06 ±0.02	83	AVERY	91	CLEO $e^+e^-$ 10.5 GeV
0.49 ±0.18 ±0.04	12	BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(pK^-\pi^+\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>4.9 ±0.4 OUR FIT</b>				Error includes scale factor of 1.1.
<b>4.53±0.23±0.30</b>	1849	ABLIKIM	16	BES3 $e^+e^- \rightarrow \Lambda_c\bar{\Lambda}_c$ , 4.599 GeV

$\Gamma(pK^-\pi^+\pi^0)/\Gamma(pK^-\pi^+)$   $\Gamma_{10}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.777±0.033 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.67 ±0.04 ±0.11</b>	2606	ALAM	98	CLE2 $e^+e^- \approx \Upsilon(4S)$

$\Gamma(pK^*(892)^-\pi^+)/\Gamma(pK_S^0\pi^+\pi^-)$   $\Gamma_{11}/\Gamma_9$   
Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.88±0.28</b>	17	ALEEV	94	BIS2 $nN$ 20–70 GeV

$\Gamma(p(K^-\pi^+)_{\text{nonresonant}}\pi^0)/\Gamma(pK^-\pi^+)$   $\Gamma_{12}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.73±0.12±0.05</b>	67	BOZEK	93	NA32 $\pi^-$ Cu 230 GeV

$\Gamma(\Delta(1232)\bar{K}^*(892))/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>seen</b>	35	AMENDOLIA	87	SPEC $\gamma$ Ge-Si

$\Gamma(pK^-\pi^+\pi^-)/\Gamma(pK^-\pi^+)$   $\Gamma_{14}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.022±0.015</b>				
		BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(pK^-\pi^+\pi^0)/\Gamma(pK^-\pi^+)$   $\Gamma_{15}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.16±0.07±0.03</b>	15	BOZEK	93	NA32 $\pi^-$ Cu 230 GeV

$\Gamma(pK^-\pi^+\pi^0)/\Gamma(pK^-\pi^+)$   $\Gamma_{16}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>• • •</b>				We do not use the following data for averages, fits, limits, etc. • • •
0.10±0.06±0.02	8	BOZEK	93	NA32 $\pi^-$ Cu 230 GeV

Hadronic modes with a  $p$ :  $S = 0$  final states

$\Gamma(p\pi^+\pi^-)/\Gamma(pK^-\pi^+)$   $\Gamma_{17}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.069±0.036</b>				
		BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(p\bar{f}_0(980))/\Gamma(pK^-\pi^+)$   $\Gamma_{18}/\Gamma_2$   
Unseen decay modes of the  $\bar{f}_0(980)$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.055±0.036</b>				
		BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(p2\pi^+2\pi^-)/\Gamma(pK^-\pi^+)$   $\Gamma_{19}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.036±0.023</b>				
		BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(pK^+K^-)/\Gamma(pK^-\pi^+)$   $\Gamma_{20}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.015±0.006 OUR AVERAGE</b>				Error includes scale factor of 2.1.
0.014±0.002±0.002	676	ABE	02c	BELL $e^+e^- \approx \Upsilon(4S)$
0.039±0.009±0.007	214	ALEXANDER	96c	CLE2 $e^+e^- \approx \Upsilon(4S)$
<b>• • •</b>				We do not use the following data for averages, fits, limits, etc. • • •
0.096±0.029±0.010	30	FRABETTI	93h	E687 $\gamma$ Be, $\bar{E}_\gamma$ 220 GeV
0.048±0.027		BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(p\phi)/\Gamma(pK^-\pi^+)$   $\Gamma_{21}/\Gamma_2$   
Unseen decay modes of the  $\phi$  are included.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.0164±0.0032 OUR AVERAGE</b>				Error includes scale factor of 1.2.
0.015 ±0.002 ±0.002	345	ABE	02c	BELL $e^+e^- \approx \Upsilon(4S)$
0.024 ±0.006 ±0.003	54	ALEXANDER	96c	CLE2 $e^+e^- \approx \Upsilon(4S)$
<b>• • •</b>				We do not use the following data for averages, fits, limits, etc. • • •
0.040 ±0.027		BARLAG	90d	NA32 $\pi^-$ 230 GeV

$\Gamma(pK^+K^-\text{non-}\phi)/\Gamma(pK^-\pi^+)$   $\Gamma_{22}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.007±0.002±0.002</b>	344	ABE	02c	BELL $e^+e^- \approx \Upsilon(4S)$

Hadronic modes with a hyperon:  $S = -1$  final states

$\Gamma(\Lambda\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{23}/\Gamma$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1.30±0.07 OUR FIT</b>				Error includes scale factor of 1.2.
<b>1.24±0.07±0.03</b>	706	ABLIKIM	16	BES3 $e^+e^- \rightarrow \Lambda_c\bar{\Lambda}_c$ , 4.599 GeV

$\Gamma(\Lambda\pi^+)/\Gamma(pK^-\pi^+)$   $\Gamma_{23}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.204±0.009 OUR FIT</b>					Error includes scale factor of 1.1.
<b>0.204±0.019 OUR AVERAGE</b>					
0.217±0.013±0.020	750	LINK	05F	FOCS	$\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV
0.18 ±0.03 ±0.04		ALBRECHT	92	ARG	$e^+e^- \approx 10.4$ GeV
0.18 ±0.03 ±0.03	87	AVERY	91	CLEO	$e^+e^-$ 10.5 GeV
<b>• • •</b>					We do not use the following data for averages, fits, limits, etc. • • •
<0.33	90	ANJOS	90	E691	$\gamma$ Be 70–260 GeV
<0.16	90	ALBRECHT	88c	ARG	$e^+e^-$ 10 GeV

$\Gamma(\Lambda\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{24}/\Gamma$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>7.1 ±0.4 OUR FIT</b>				Error includes scale factor of 1.2.
<b>7.01±0.37±0.19</b>	1497	ABLIKIM	16	BES3 $e^+e^- \rightarrow \Lambda_c\bar{\Lambda}_c$ , 4.599 GeV

$\Gamma(\Lambda\pi^+\pi^0)/\Gamma(pK^-\pi^+)$   $\Gamma_{24}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>1.11±0.05 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.73±0.09±0.16</b>	464	AVERY	94	CLE2 $e^+e^- \approx \Upsilon(3S), \Upsilon(4S)$

$\Gamma(\Lambda\rho^+)/\Gamma(pK^-\pi^+)$   $\Gamma_{25}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.95</b>	95	AVERY	94	CLE2 $e^+e^- \approx \Upsilon(3S), \Upsilon(4S)$

$\Gamma(\Lambda\pi^-2\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{26}/\Gamma$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE (%)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>3.7 ±0.4 OUR FIT</b>				Error includes scale factor of 1.9.
<b>3.81±0.24±0.18</b>	609	ABLIKIM	16	BES3 $e^+e^- \rightarrow \Lambda_c\bar{\Lambda}_c$ , 4.599 GeV

$\Gamma(\Lambda\pi^-2\pi^+)/\Gamma(pK^-\pi^+)$   $\Gamma_{26}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.58 ±0.06 OUR FIT</b>				Error includes scale factor of 2.8.
<b>0.522±0.032 OUR AVERAGE</b>				
0.508±0.024±0.024	1356	LINK	05F	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV
0.65 ±0.11 ±0.12	289	AVERY	91	CLEO $e^+e^-$ 10.5 GeV
0.82 ±0.29 ±0.27	44	ANJOS	90	E691 $\gamma$ Be 70–260 GeV
0.94 ±0.41 ±0.13	10	BARLAG	90d	NA32 $\pi^-$ 230 GeV
0.61 ±0.16 ±0.04	105	ALBRECHT	88c	ARG $e^+e^-$ 10 GeV

$\Gamma(\Sigma(1385)^+\pi^+\pi^-, \Sigma^{*+} \rightarrow \Lambda\pi^+)/\Gamma(\Lambda\pi^-2\pi^+)$   $\Gamma_{27}/\Gamma_{26}$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.28±0.10±0.08</b>	LINK	05F	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\Sigma(1385)^-\pi^+\pi^-, \Sigma^{*-} \rightarrow \Lambda\pi^-)/\Gamma(\Lambda\pi^-2\pi^+)$   $\Gamma_{28}/\Gamma_{26}$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.21±0.03±0.02</b>	LINK	05F	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\Lambda\pi^+\rho^0)/\Gamma(\Lambda\pi^-2\pi^+)$   $\Gamma_{29}/\Gamma_{26}$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.40±0.12±0.12</b>	LINK	05F	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\Sigma(1385)^+\rho^0, \Sigma^{*+} \rightarrow \Lambda\pi^+)/\Gamma(\Lambda\pi^-2\pi^+)$   $\Gamma_{30}/\Gamma_{26}$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.14±0.09±0.07</b>	LINK	05F	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\Lambda\pi^-2\pi^+\text{nonresonant})/\Gamma(\Lambda\pi^-2\pi^+)$   $\Gamma_{31}/\Gamma_{26}$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.3</b>	90	LINK	05F	FOCS $\gamma$ nucleus, $\bar{E}_\gamma \approx 180$ GeV

$\Gamma(\Lambda\pi^-\pi^02\pi^+\text{total})/\Gamma(pK^-\pi^+)$   $\Gamma_{32}/\Gamma_2$   
Measurements given as a  $\bar{K}^0$  ratio have been divided by 2 to convert to a  $K_S^0$  ratio.

VALUE	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.36±0.09±0.09</b>	50	1	CRONIN-HENNESSY 03	CLE3 $e^+e^- \approx \Upsilon(4S)$

<sup>1</sup> CRONIN-HENNESSY 03 finds this channel to be dominantly  $\Lambda\eta\pi^+$  and  $\Lambda\omega\pi^+$ ; see below.

# Baryon Particle Listings

$\Lambda^+_C$

$\Gamma(\Lambda\pi^+\eta)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{33}/\Gamma_2</math></span>				
Unseen decay modes of the $\eta$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.36±0.07 OUR AVERAGE</b>				
0.41±0.17±0.10	11	CRONIN-HEN..03	CLE3	$e^+e^-\approx\mathcal{T}(45)$
0.35±0.05±0.06	116	AMMAR	95	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Sigma(1385)^+\eta)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{34}/\Gamma_2</math></span>				
Unseen decay modes of the $\Sigma(1385)^+$ and $\eta$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.17±0.04±0.03</b>	54	AMMAR	95	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Lambda\pi^+\omega)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{35}/\Gamma_2</math></span>				
Unseen decay modes of the $\omega$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.24±0.06±0.06</b>	32	CRONIN-HEN..03	CLE3	$e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Lambda\pi^-\pi^0\,2\pi^+, \text{ no } \eta \text{ or } \omega)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{36}/\Gamma_2</math></span>				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.13</b>	90	CRONIN-HEN..03	CLE3	$e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Lambda K^+\bar{K}^0)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{37}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.089±0.018 OUR FIT</b>	Error includes scale factor of 2.0.			
<b>0.131±0.020 OUR AVERAGE</b>				
0.142±0.018±0.022	251	LINK	05F	FOCS $\gamma$ nucleus, $\overline{E}_\gamma\approx 180$ GeV
0.12±0.02±0.02	59	AMMAR	95	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Xi(1690)^0K^+,\Xi^{*0}\rightarrow\Lambda\bar{K}^0)/\Gamma(\Lambda K^+\bar{K}^0)$ <span style="float:right"><math>\Gamma_{38}/\Gamma_{37}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.28±0.07 OUR AVERAGE</b>				
0.32±0.10±0.04	84±24	LINK	05F	FOCS $\gamma$ nucleus, $\overline{E}_\gamma\approx 180$ GeV
0.26±0.08±0.03	93	ABE	02c	BELL $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Lambda K^+\bar{K}^0)/\Gamma(\Lambda\pi^+)$ <span style="float:right"><math>\Gamma_{37}/\Gamma_{23}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.44±0.08 OUR FIT</b>	Error includes scale factor of 2.1.			
<b>0.395±0.026±0.036</b>	460±30	AUBERT	07u	BABR $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Sigma^0\pi^+)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{39}/\Gamma</math></span>				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.29±0.07 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>1.27±0.08±0.03</b>	522	ABLIKIM	16	BES3 $e^+e^-\rightarrow\Lambda_C\bar{\Lambda}_C$ , 4.599 GeV

$\Gamma(\Sigma^0\pi^+)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{39}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.203±0.010 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.20±0.04 OUR AVERAGE</b>				
0.21±0.02±0.04	196	AVERY	94	CLE2 $e^+e^-\approx\mathcal{T}(35),\mathcal{T}(45)$
0.17±0.06±0.04		ALBRECHT	92	ARG $e^+e^-\approx 10.4$ GeV

$\Gamma(\Sigma^0\pi^+)/\Gamma(\Lambda\pi^+)$ <span style="float:right"><math>\Gamma_{39}/\Gamma_{23}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.99±0.04 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>0.98±0.05 OUR AVERAGE</b>				
0.977±0.015±0.051	33k	AUBERT	07u	BABR $e^+e^-\approx\mathcal{T}(45)$
1.09±0.11±0.19	750	LINK	05F	FOCS $\gamma$ nucleus, $\overline{E}_\gamma\approx 180$ GeV

$\Gamma(\Sigma^+\pi^0)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{40}/\Gamma</math></span>				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.24±0.10 OUR FIT</b>				
<b>1.18±0.10±0.03</b>	309	ABLIKIM	16	BES3 $e^+e^-\rightarrow\Lambda_C\bar{\Lambda}_C$ , 4.599 GeV

$\Gamma(\Sigma^+\pi^0)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{40}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.196±0.015 OUR FIT</b>				
<b>0.20±0.03±0.03</b>	93	KUBOTA	93	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Sigma^+\eta)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{41}/\Gamma_2</math></span>				
Unseen decay modes of the $\eta$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.11±0.03±0.02</b>	26	AMMAR	95	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Sigma^+\pi^+\pi^-)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{42}/\Gamma</math></span>				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.57±0.29 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>4.25±0.24±0.20</b>	1156	ABLIKIM	16	BES3 $e^+e^-\rightarrow\Lambda_C\bar{\Lambda}_C$ , 4.599 GeV

$\Gamma(\Sigma^+\pi^+\pi^-)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{42}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.720±0.029 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>0.69±0.08 OUR AVERAGE</b>				
0.72±0.14	47±9	VAZQUEZ-JA...08	SELX	$\Sigma^-$ nucleus, 600 GeV
0.74±0.07±0.09	487	KUBOTA	93	CLE2 $e^+e^-\approx\mathcal{T}(45)$
0.54+0.18-0.15	11	BARLAG	92	NA32 $\pi^-$ Cu 230 GeV

$\Gamma(\Sigma^+\rho^0)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{43}/\Gamma_2</math></span>				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.27</b>	95	KUBOTA	93	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Sigma^-2\pi^+)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{44}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.33±0.06 OUR FIT</b>	30±6	VAZQUEZ-JA...08	SELX	$\Sigma^-$ nucleus, 600 GeV
<b>0.314±0.067</b>				

$\Gamma(\Sigma^-2\pi^+)/\Gamma(\Sigma^+\pi^+\pi^-)$ <span style="float:right"><math>\Gamma_{44}/\Gamma_{42}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.46±0.08 OUR FIT</b>				
<b>0.53±0.15±0.07</b>	56	FRABETTI	94E	E687 $\gamma$ Be, $\overline{E}_\gamma$ 220 GeV

$\Gamma(\Sigma^0\pi^+\pi^0)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{45}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.36±0.09±0.10</b>	117	AVERY	94	CLE2 $e^+e^-\approx\mathcal{T}(35),\mathcal{T}(45)$

$\Gamma(\Sigma^0\pi^-\,2\pi^+)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{46}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.18±0.05 OUR FIT</b>				
<b>0.21±0.05±0.05</b>	90	AVERY	94	CLE2 $e^+e^-\approx\mathcal{T}(35),\mathcal{T}(45)$

$\Gamma(\Sigma^0\pi^-\,2\pi^+)/\Gamma(\Lambda\pi^-\,2\pi^+)$ <span style="float:right"><math>\Gamma_{46}/\Gamma_{26}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.30±0.08 OUR FIT</b>				
<b>0.26±0.06±0.09</b>	480	LINK	05F	FOCS $\gamma$ nucleus, $\overline{E}_\gamma\approx 180$ GeV

$\Gamma(\Sigma^+\omega)/\Gamma_{\text{total}}$ <span style="float:right"><math>\Gamma_{48}/\Gamma</math></span>				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.74±0.21 OUR FIT</b>				
<b>1.56±0.20±0.07</b>	157	ABLIKIM	16	BES3 $e^+e^-\rightarrow\Lambda_C\bar{\Lambda}_C$ , 4.599 GeV

$\Gamma(\Sigma^+\omega)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{48}/\Gamma_2</math></span>				
Unseen decay modes of the $\omega$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.274±0.032 OUR FIT</b>				
<b>0.54±0.13±0.06</b>	107	KUBOTA	93	CLE2 $e^+e^-\approx\mathcal{T}(45)$

$\Gamma(\Sigma^+K^+K^-)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{49}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.056±0.006 OUR FIT</b>				
<b>0.070±0.011±0.011</b>	59	AVERY	93	CLE2 $e^+e^-\approx 10.5$ GeV

$\Gamma(\Sigma^+K^+K^-)/\Gamma(\Sigma^+\pi^+\pi^-)$ <span style="float:right"><math>\Gamma_{49}/\Gamma_{42}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.078±0.008 OUR FIT</b>				
<b>0.074±0.009 OUR AVERAGE</b>				
0.076±0.007±0.009	246	ABE	02c	BELL $e^+e^-\approx\mathcal{T}(45)$
0.071±0.011±0.011	103	LINK	02G	FOCS $\gamma$ nucleus, $\approx 180$ GeV

$\Gamma(\Sigma^+\phi)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{50}/\Gamma_2</math></span>				
Unseen decay modes of the $\phi$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.063±0.009 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>0.069±0.023±0.016</b>	26	AVERY	93	CLE2 $e^+e^-\approx 10.5$ GeV

$\Gamma(\Sigma^+\phi)/\Gamma(\Sigma^+\pi^+\pi^-)$ <span style="float:right"><math>\Gamma_{50}/\Gamma_{42}</math></span>				
Unseen decay modes of the $\phi$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.087±0.012 OUR FIT</b>				
<b>0.086±0.012 OUR AVERAGE</b>				
0.085±0.012±0.012	129	ABE	02c	BELL $e^+e^-\approx\mathcal{T}(45)$
0.087±0.016±0.006	57	LINK	02G	FOCS $\gamma$ nucleus, $\approx 180$ GeV

$\Gamma(\Xi(1690)^0K^+,\Xi^{*0}\rightarrow\Sigma^+K^-)/\Gamma(\Sigma^+\pi^+\pi^-)$ <span style="float:right"><math>\Gamma_{51}/\Gamma_{42}</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.023±0.005 OUR AVERAGE</b>				
0.023±0.005±0.005	75	ABE	02c	BELL $e^+e^-\approx\mathcal{T}(45)$
0.022±0.006±0.006	34	LINK	02G	FOCS $\gamma$ nucleus, $\approx 180$ GeV

$\Gamma(\Sigma^+K^+K^-\text{ nonresonant})/\Gamma(\Sigma^+\pi^+\pi^-)$ <span style="float:right"><math>\Gamma_{52}/\Gamma_{42}</math></span>				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.018</b>	90	ABE	02c	BELL $e^+e^-\approx\mathcal{T}(45)$
••• We do not use the following data for averages, fits, limits, etc. •••				
<0.028	90	LINK	02G	FOCS $\gamma$ nucleus, $\approx 180$ GeV

$\Gamma(\Xi^0K^+)/\Gamma(pK^-\pi^+)$ <span style="float:right"><math>\Gamma_{53}/\Gamma_2</math></span>				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.078±0.013±0.013</b>	56	AVERY	93	CLE2 $e^+e^-\approx 10.5$ GeV



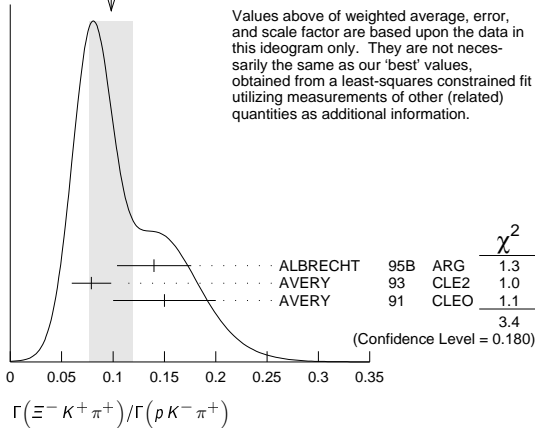
See key on page 601

## Baryon Particle Listings

 $\Lambda_c^+$ 

$\Gamma(\Xi^- K^+ \pi^+)/\Gamma(p K^- \pi^+)$		$\Gamma_{54}/\Gamma_2$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.098±0.009 OUR FIT</b>	Error	includes scale factor of 1.1.	
<b>0.098±0.021 OUR AVERAGE</b>	Error	includes scale factor of 1.3. See the ideogram below.	
0.14 ±0.03 ±0.02	34	ALBRECHT	95B ARG $e^+ e^- \approx 10.4$ GeV
0.079±0.013±0.014	60	AVERY	93 CLE2 $e^+ e^- \approx 10.5$ GeV
0.15 ±0.04 ±0.03	30	AVERY	91 CLEO $e^+ e^- 10.5$ GeV

WEIGHTED AVERAGE  
0.098±0.021 (Error scaled by 1.3)



$\Gamma(\Xi(1530)^0 K^+)/\Gamma(p K^- \pi^+)$				$\Gamma_{55}/\Gamma_2$
Unseen decay modes of the $\Xi(1530)^0$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.052±0.014 OUR AVERAGE</b>				
0.05 ±0.02 ±0.01	11	ALBRECHT	95B ARG	$e^+ e^- \approx 10.4$ GeV
0.053±0.016±0.010	24	AVERY	93 CLE2	$e^+ e^- \approx 10.5$ GeV

$\Gamma(\Xi^- K^+ \pi^+)/\Gamma(\Lambda \pi^+)$		$\Gamma_{54}/\Gamma_{23}$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.48 ±0.04 OUR FIT</b>			
<b>0.480±0.016±0.039</b>	2665 ±84	AUBERT	07u BABR $e^+ e^- \approx \gamma(4S)$

Hadronic modes with a hyperon:  $S = 0$  final states

$\Gamma(\Lambda K^+)/\Gamma(\Lambda \pi^+)$		$\Gamma_{56}/\Gamma_{23}$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.047±0.009 OUR AVERAGE</b>	Error	includes scale factor of 1.8.	
0.044±0.004±0.003	1162 ±101	AUBERT	07u BABR $e^+ e^- \approx \gamma(4S)$
0.074±0.010±0.012	265	ABE	02c BELL $e^+ e^- \approx \gamma(4S)$

$\Gamma(\Lambda K^+ \pi^-)/\Gamma(\Lambda \pi^+)$		$\Gamma_{57}/\Gamma_{23}$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;4.1 × 10<sup>-2</sup></b>	90	AUBERT	07u BABR $e^+ e^- \approx \gamma(4S)$

$\Gamma(\Sigma^0 K^+)/\Gamma(\Sigma^0 \pi^+)$		$\Gamma_{58}/\Gamma_{39}$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.040±0.006 OUR AVERAGE</b>			
0.038±0.005±0.003	366 ±52	AUBERT	07u BABR $e^+ e^- \approx \gamma(4S)$
0.056±0.014±0.008	75	ABE	02c BELL $e^+ e^- \approx \gamma(4S)$

$\Gamma(\Sigma^0 K^+ \pi^-)/\Gamma(\Sigma^0 \pi^-)$		$\Gamma_{59}/\Gamma_{39}$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;2.0 × 10<sup>-2</sup></b>	90	AUBERT	07u BABR $e^+ e^- \approx \gamma(4S)$

$\Gamma(\Sigma^+ K^+ \pi^-)/\Gamma(\Sigma^+ \pi^+ \pi^-)$		$\Gamma_{60}/\Gamma_{42}$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.047±0.011±0.008</b>	105	ABE	02c BELL $e^+ e^- \approx \gamma(4S)$

$\Gamma(\Sigma^+ K^*(892)^0)/\Gamma(\Sigma^+ \pi^+ \pi^-)$	$\Gamma_{61}/\Gamma_{42}$
Unseen decay modes of the $K^*(892)^0$ are included.	
VALUE	EVTS
<b>0.078±0.018±0.013</b>	49

$\Gamma(\Sigma^- K^+ \pi^+)/\Gamma(\Sigma^+ K^*(892)^0)$		$\Gamma_{62}/\Gamma_{61}$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.35</b>	90	LINK	02G FOCS $\gamma$ nucleus, $\approx 180$ GeV

## Doubly Cabibbo-suppressed modes

$\Gamma(p K^+ \pi^-)/\Gamma(p K^- \pi^+)$		$\Gamma_{63}/\Gamma_2$	
VALUE	CL%	DOCUMENT ID	TECN COMMENT
<b>&lt;0.0046</b>	90	LINK	05K FOCS $R = (0.05 \pm 0.26 \pm 0.02)\%$

## Semileptonic modes

$\Gamma(\Lambda e^+ \nu_e)/\Gamma_{\text{total}}$		$\Gamma_{64}/\Gamma$	
VALUE (%)	EVTS	DOCUMENT ID	TECN COMMENT
<b>3.63±0.38±0.20</b>	104	ABLIKIM	15Y BES3 567 pb <sup>-1</sup> , 4.599 GeV

$\Gamma(\Lambda e^+ \nu_e)/\Gamma(p K^- \pi^+)$		$\Gamma_{64}/\Gamma_2$	
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.43±0.08	<sup>1,2</sup> BERGFELD 94	CLE2	$e^+ e^- \approx \gamma(4S)$
0.38±0.14	<sup>2,3</sup> ALBRECHT 91G	ARG	$e^+ e^- \approx 10.4$ GeV

<sup>1</sup> BERGFELD 94 measures  $\sigma(e^+ e^- \rightarrow \Lambda_c^+ X) \cdot B(\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e) = (4.87 \pm 0.28 \pm 0.69)$  pb.

<sup>2</sup> To extract  $\Gamma(\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e)/\Gamma(\Lambda_c^+ \rightarrow p K^- \pi^+)$ , we use  $\sigma(e^+ e^- \rightarrow \Lambda_c^+ X) \cdot B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (11.2 \pm 1.3)$  pb, which is the weighted average of measurements from ARGUS (ALBRECHT 96E) and CLEO (AVERY 91).

<sup>3</sup> ALBRECHT 91G measures  $\sigma(e^+ e^- \rightarrow \Lambda_c^+ X) \cdot B(\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e) = (4.20 \pm 1.28 \pm 0.71)$  pb.

$\Gamma(\Lambda \mu^+ \nu_\mu)/\Gamma(p K^- \pi^+)$		$\Gamma_{65}/\Gamma_2$	
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.40±0.09	<sup>1,2</sup> BERGFELD 94	CLE2	$e^+ e^- \approx \gamma(4S)$
0.35±0.20	<sup>2,3</sup> ALBRECHT 91G	ARG	$e^+ e^- \approx 10.4$ GeV

<sup>1</sup> BERGFELD 94 measures  $\sigma(e^+ e^- \rightarrow \Lambda_c^+ X) \cdot B(\Lambda_c^+ \rightarrow \Lambda \mu^+ \nu_\mu) = (4.43 \pm 0.51 \pm 0.64)$  pb.

<sup>2</sup> To extract  $\Gamma(\Lambda_c^+ \rightarrow \Lambda \mu^+ \nu_\mu)/\Gamma(\Lambda_c^+ \rightarrow p K^- \pi^+)$ , we use  $\sigma(e^+ e^- \rightarrow \Lambda_c^+ X) \cdot B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (11.2 \pm 1.3)$  pb, which is the weighted average of measurements from ARGUS (ALBRECHT 96E) and CLEO (AVERY 91).

<sup>3</sup> ALBRECHT 91G measures  $\sigma(e^+ e^- \rightarrow \Lambda_c^+ X) \cdot B(\Lambda_c^+ \rightarrow \Lambda \mu^+ \nu_\mu) = (3.91 \pm 2.02 \pm 0.90)$  pb.

## Inclusive modes

$\Gamma(e^+ \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{66}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.045±0.017</b>	VELLA	82	MRK2 $e^+ e^- 4.5\text{--}6.8$ GeV

$\Gamma(p e^+ \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{67}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.018±0.009</b>	<sup>1</sup> VELLA	82	MRK2 $e^+ e^- 4.5\text{--}6.8$ GeV
<sup>1</sup> VELLA 82 includes protons from $\Lambda$ decay.			

$\Gamma(\Lambda e^+ \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{68}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.011±0.008	<sup>1</sup> VELLA	82	MRK2 $e^+ e^- 4.5\text{--}6.8$ GeV
<sup>1</sup> VELLA 82 includes $\Lambda$ 's from $\Sigma^0$ decay.			

$\Gamma(p \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{69}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.50±0.08±0.14</b>	<sup>1</sup> CRAWFORD 92	CLEO	$e^+ e^- 10.5$ GeV
<sup>1</sup> This CRAWFORD 92 value includes protons from $\Lambda$ decay. The value is model dependent, but account is taken of this in the systematic error.			

$\Gamma(p \text{ anything (no } \Lambda))/\Gamma_{\text{total}}$		$\Gamma_{70}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.12±0.10±0.16</b>	CRAWFORD 92	CLEO	$e^+ e^- 10.5$ GeV

$\Gamma(n \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{72}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.50±0.08±0.14</b>	<sup>1</sup> CRAWFORD 92	CLEO	$e^+ e^- 10.5$ GeV
<sup>1</sup> This CRAWFORD 92 value includes neutrons from $\Lambda$ decay. The value is model dependent, but account is taken of this in the systematic error.			

$\Gamma(n \text{ anything (no } \Lambda))/\Gamma_{\text{total}}$		$\Gamma_{73}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.29±0.09±0.15</b>	CRAWFORD 92	CLEO	$e^+ e^- 10.5$ GeV

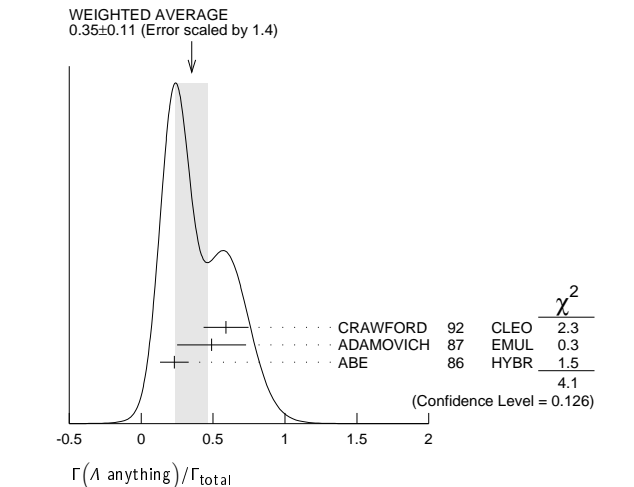
$\Gamma(p \text{ hadrons})/\Gamma_{\text{total}}$		$\Gamma_{71}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.41±0.24	ADAMOVICH 87	EMUL	$\gamma A$ 20-70 GeV/c

$\Gamma(\Lambda \text{ anything})/\Gamma_{\text{total}}$		$\Gamma_{74}/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN COMMENT
<b>0.35±0.11 OUR AVERAGE</b>	Error	includes scale factor of 1.4. See the ideogram below.	
0.59±0.10±0.12		CRAWFORD 92	CLEO $e^+ e^- 10.5$ GeV
0.49±0.24		ADAMOVICH 87	EMUL $\gamma A$ 20-70 GeV/c
0.23±0.10	8	<sup>1</sup> ABE 86	HYBR 20 GeV $\gamma p$

# Baryon Particle Listings

$\Lambda_c^+$

<sup>1</sup> ABE 86 includes  $\Lambda$ 's from  $\Sigma^0$  decay.



$\Gamma(\Sigma^\pm \text{ anything})/\Gamma_{\text{total}}$				$\Gamma_{75}/\Gamma$	
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
0.1±0.05		5	ABE 86	HYBR	20 GeV $\gamma p$

$\Gamma(3\text{prongs})/\Gamma_{\text{total}}$				$\Gamma_{76}/\Gamma$	
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
0.24±0.07±0.04			KAYIS-TOPAK.03	CHRS	$\nu_\mu$ emulsion, $\bar{E}=27$ GeV

## Rare or forbidden modes

$\Gamma(p e^+ e^-)/\Gamma_{\text{total}}$				$\Gamma_{77}/\Gamma$	
A test for the $\Delta C=1$ weak neutral current. Allowed by higher-order electroweak interactions.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<5.5 \times 10^{-6}$	90	$4.0 \pm 7.1$	LEES	11G	BABR $e^+e^- \approx \Upsilon(4S)$

$\Gamma(p\mu^+\mu^-)/\Gamma_{\text{Total}}$				$\Gamma_{78}/\Gamma$	
A test for the $\Delta C=1$ weak neutral current. Allowed by higher-order electroweak interactions.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;44 × 10<sup>-6</sup></b>	90	11.1 ± 5.6	LEES	11G	BABR $e^+e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 3.4 × 10 <sup>-4</sup>	90	0	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV

$\Gamma(p e^+ \mu^-)/\Gamma_{\text{total}}$					$\Gamma_{79}/\Gamma$
A test of lepton family-number conservation.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<9.9 \times 10^{-6}$	90	$-0.7 \pm 3.0$	LEES	11G BABR	$e^+ e^- \approx \Upsilon(4S)$

$\Gamma(p e^- \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{80}/\Gamma$
A test of lepton family-number conservation.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<19 \times 10^{-6}$	90	$6.2 \pm 4.9$	LEES	11G BABR	$e^+ e^- \approx \Upsilon(4S)$

$\Gamma(\bar{p}2e^+)/\Gamma_{\text{total}}$				$\Gamma_{81}/\Gamma$	
A test of lepton- and baryon-number conservation.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.7 \times 10^{-6}$	90	$-1.5 \pm 4.5$	LEES	11G	BABR $e^+e^- \approx \Upsilon(4S)$

$\Gamma(\bar{p}2\mu^+)/\Gamma_{\text{total}}$				$\Gamma_{82}/\Gamma$	
A test of lepton- and baryon-number conservation and of lepton family-number conservation.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<9.4 \times 10^{-6}$	90	$0.0 \pm 2.2$	LEES	11G	BABR $e^+e^- \approx \Upsilon(4S)$

$\Gamma(\bar{p}e^+\mu^+)/\Gamma_{\text{total}}$					$\Gamma_{83}/\Gamma$
A test of lepton- and baryon-number conservation and of lepton family-number conservation.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<16 \times 10^{-6}$	90	$10.1 \pm 6.8$	LEES	11G BABR	$e^+e^- \approx \Upsilon(4S)$

$\Gamma(\Sigma^- \mu^+ \mu^+)/\Gamma_{\text{total}}$					$\Gamma_{84}/\Gamma$
A test of lepton-number conservation.					
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<7.0 \times 10^{-4}$	90	0	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

## $\Lambda_c^+$ DECAY PARAMETERS

See the note on "Baryon Decay Parameters" in the neutron Listings.

$\alpha$ FOR $\Lambda_c^+ \rightarrow \Lambda \pi^+$			
VALUE	EVTS	DOCUMENT ID	TECN
-0.91±0.15 OUR AVERAGE			
-0.78±0.16±0.19		LINK	06A
-0.94±0.21±0.12	414	<sup>1</sup> BISHAI	95
-0.96±0.42		ALBRECHT	92
-1.1 ± 0.4	86	AVERY	90B

<sup>1</sup> BISHAI 95 actually gives  $\alpha = -0.94 \pm 0.21 \pm 0.12$ , chopping the errors at the physical limit  $-1.0$ . However, for  $\alpha \approx -1.0$ , some experiments should get unphysical values ( $\alpha < -1.0$ ), and for averaging with other measurements such values (or errors that extend below  $-1.0$ ) should not be chopped.

$\alpha$ FOR $\Lambda_c^+ \rightarrow \Sigma^+ \pi^0$			
VALUE	EVTS	DOCUMENT ID	TECN
-0.45±0.31±0.06	89	BISHAI	95

$\alpha$ FOR $\Lambda_c^+ \rightarrow \Lambda \ell^+ \nu_\ell$			
The experiments don't cover the complete (or same incomplete) $M(\Lambda \ell^+)$ range, but we average them together anyway.			
VALUE	EVTS	DOCUMENT ID	TECN
-0.86±0.04 OUR AVERAGE			
-0.86±0.03±0.02	3201	<sup>1</sup> HINSON	05
-0.91±0.42±0.25		<sup>2</sup> ALBRECHT	94B
-0.82±0.09±0.06	700	<sup>3</sup> CRAWFORD	95
-0.89±0.17±0.09	350	<sup>4</sup> BERGFELD	94

- • • We do not use the following data for averages, fits, limits, etc. • • •
- <sup>1</sup> HINSON 05 measures the form-factor ratio  $R \equiv f_2/f_1$  for  $\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e$  events to be  $-0.31 \pm 0.05 \pm 0.04$  and the pole mass to be  $2.21 \pm 0.08 \pm 0.14$  GeV/c<sup>2</sup>, and from these calculates  $\alpha$ , averaged over  $q^2$ , where  $\langle q^2 \rangle = 0.67$  (GeV/c)<sup>2</sup>.
- <sup>2</sup> ALBRECHT 94B uses  $\Lambda e^+$  and  $\Lambda \mu^+$  events in the mass range  $1.85 < M(\Lambda \ell^+) < 2.20$  GeV.
- <sup>3</sup> CRAWFORD 95 measures the form-factor ratio  $R \equiv f_2/f_1$  for  $\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e$  events to be  $-0.25 \pm 0.14 \pm 0.08$  and from this calculates  $\alpha$ , averaged over  $q^2$ , to be the above.
- <sup>4</sup> BERGFELD 94 uses  $\Lambda e^+$  events.

## $\Lambda_c^+, \bar{\Lambda}_c^-$ CP-VIOLATING DECAY ASYMMETRIES

$(\alpha + \bar{\alpha})/(\alpha - \bar{\alpha})$ in $\Lambda_c^+ \rightarrow \Lambda \pi^+, \bar{\Lambda}_c^- \rightarrow \bar{\Lambda} \pi^-$			
This is zero if CP is conserved.			
VALUE	DOCUMENT ID	TECN	COMMENT
-0.07±0.19±0.24	LINK	06A	FOCS $\gamma A, \bar{E}_\gamma \approx 180$ GeV

$(\alpha + \bar{\alpha})/(\alpha - \bar{\alpha})$ in $\Lambda_c^+ \rightarrow \Lambda e^+ \nu_e, \bar{\Lambda}_c^- \rightarrow \bar{\Lambda} e^- \bar{\nu}_e$			
This is zero if CP is conserved.			
VALUE	DOCUMENT ID	TECN	COMMENT
0.00±0.03±0.02	HINSON	05	CLEO $e^+e^- \approx \gamma(4S)$

## $\Lambda_c^+$ REFERENCES

We have omitted some papers that have been superseded by later experiments. The omitted papers may be found in our 1992 edition (Physical Review D45, 1 June, Part II) or in earlier editions.

ABLIKIM	16	PRL 116 052001	M. Ablikim et al.	(BES III Collab.)
ABLIKIM	15Y	PRL 115 221805	M. Ablikim et al.	(BES III Collab.)
ZUPANC	14	PRL 113 042002	A. Zupanc et al.	(BELLE Collab.)
LEES	11G	PR D84 072006	J.P. Lees et al.	(BABAR Collab.)
VAZQUEZ-JA...	08	PL B666 299	E. Vazquez-Jauregui et al.	(SELEX Collab.)
AUBERT	07U	PR D75 052002	B. Aubert et al.	(BABAR Collab.)
LINK	06A	PL B634 165	J.M. Link et al.	(FNAL FOCUS Collab.)
AUBERT,B	05S	PR D72 052006	B. Aubert et al.	(BABAR Collab.)
HINSON	05	PRL 94 191801	J.W. Hinson et al.	(CLEO Collab.)
LINK	05F	PL B624 22	J.M. Link et al.	(FNAL FOCUS Collab.)
LINK	05K	PL B624 166	J.M. Link et al.	(FNAL FOCUS Collab.)
CRONIN-HEN...	03	PR D67 012001	D. Cronin-Hennessy et al.	(CLEO Collab.)
KAYIS-TOPAK.	03	PL B555 156	A. Kayis-Topaksu et al.	(CERN CHORUS Collab.)
ABE	02C	PL B524 33	K. Abe et al.	(KEK BELLE Collab.)
LINK	02C	PRL 88 161801	J.M. Link et al.	(FNAL FOCUS Collab.)
LINK	02G	PL B540 25	J.M. Link et al.	(FNAL FOCUS Collab.)
PDG	02	PR D66 010001	K. Hagiwara et al.	(PDG Collab.)
KUSHNIR...	01	PRL 86 5243	A. Kushnirenko et al.	(FNAL SELEX Collab.)
MAHMOOD	01	PRL 86 2232	A.H. Mahmood et al.	(CLEO Collab.)
AITALA	00	PL B471 449	E.M. Aitala et al.	(FNAL E791 Collab.)
ALAM	98	PR D57 4467	M.S. Alam et al.	(CLEO Collab.)
ALBRECHT	96E	PRPL 276 223	H. Albrecht et al.	(ARGUS Collab.)
ALEXANDER	96C	PR D53 R1013	J.P. Alexander et al.	(CLEO Collab.)
ALBRECHT	95B	PL B342 397	H. Albrecht et al.	(ARGUS Collab.)
AMMAR	95	PRL 74 3534	R. Ammar et al.	(CLEO Collab.)
BISHAI	95	PL B350 256	M. Bishai et al.	(CLEO Collab.)
CRAWFORD	95	PRL 75 624	G. Crawford et al.	(CLEO Collab.)
KODAMA	95	PL B345 85	K. Kodama et al.	(FNAL E653 Collab.)
ALBRECHT	94B	PL B326 320	H. Albrecht et al.	(ARGUS Collab.)
ALEEV	94	PAN 57 1370	A.N. Aleev et al.	(Serpukhov BIS-2 Collab.)

Translated from YF 57 1443.

See key on page 601

## Baryon Particle Listings

$$\Lambda_c^+, \Lambda_c(2595)^+, \Lambda_c(2625)^+$$

AVERY	94	PL B325 257	P. Avery <i>et al.</i>	(CLEO Collab.)
BERGFELD	94	PL B323 219	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
FRABETTI	94E	PL B328 193	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
AVERY	93	PRL 71 2391	P. Avery <i>et al.</i>	(CLEO Collab.)
BOZEK	93	PL B312 247	A. Bozek <i>et al.</i>	(CERN NA32 Collab.)
FRABETTI	93D	PRL 70 1755	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	93H	PL B314 477	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KUBOTA	93	PRL 71 3255	Y. Kubota <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92	PL B274 239	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARLAG	92	PL B283 465	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
CRAWFORD	92	PR D45 752	G. Crawford <i>et al.</i>	(CLEO Collab.)
JEZABEK	92	PL B286 175	M. Jezabek, K. Rybicki, R. Rylko	(CRAC)
ALBRECHT	91G	PL B269 234	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AVERY	91	PR D43 3599	P. Avery <i>et al.</i>	(CLEO Collab.)
ALVAREZ	90	ZPHY C47 539	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
ALVAREZ	90B	PL B246 256	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
ANJOS	90	PR D41 801	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
AVERY	90B	PRL 65 2842	P. Avery <i>et al.</i>	(CLEO Collab.)
BARLAG	90D	ZPHY C48 29	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
FRABETTI	90	PL B251 639	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
BARLAG	89	PL B218 374	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
AGUILAR...	88B	ZPHY C40 321	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also		PL B189 254	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also		PL B199 462	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also		SJNP 48 833	M. Begalli <i>et al.</i>	(LEBC-EHS Collab.)
Translated from YAF 48 1310.				
ALBRECHT	88C	PL B207 109	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	88B	PRL 60 1379	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ADAMOVICH	87	EPL 4 887	M.I. Adamovich <i>et al.</i>	(Photon Emulsion Collab.)
Also		SJNP 46 447	F. Viaggi <i>et al.</i>	(Photon Emulsion Collab.)
Translated from YAF 46 793				
AMENDOLIA	87	ZPHY C36 513	S.R. Amendolia <i>et al.</i>	(CERN NA1 Collab.)
JONES	87	ZPHY C36 593	G.T. Jones <i>et al.</i>	(CERN WA21 Collab.)
ABE	86	PR D33 1	K. Abe <i>et al.</i>	(SLAC HF Photon Collab.)
BOSETTI	82	PL 109B 234	P.C. Bosetti <i>et al.</i>	(AACH3, BONN, CERN+)
VELLA	82	PRL 48 1515	E. Vella <i>et al.</i>	(SLAC, LBL, UCB)
BASILE	81B	NC 62A 14	M. Basile <i>et al.</i>	(CERN, BGNA, PGIA, FRAS)
CALICCHIO	80	PL 93B 521	M. Calicchio <i>et al.</i>	(BARI, BIRM, BRUX+)

## OTHER RELATED PAPERS

MIGLIOZZI	99	PL B462 217	P. Migliozi <i>et al.</i>
DUNIETZ	98	PR D58 094010	I. Dunietz

$$\Lambda_c(2595)^+$$

$$J(P) = 0(\frac{1}{2}^-) \text{ Status: } ***$$

The  $\Lambda_c^+ \pi^+ \pi^-$  mode is largely, and perhaps entirely,  $\Sigma_c \pi$ , which is just at threshold; since the  $\Sigma_c$  has  $J^P = 1/2^+$ , the  $J^P$  here is almost certainly  $1/2^-$ . This result is in accord with the theoretical expectation that this is the charm counterpart of the strange  $\Lambda(1405)$ .

 $\Lambda_c(2595)^+$  MASS

The mass is obtained from the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass-difference measurements below.

VALUE (MeV)	DOCUMENT ID
<b>2592.25 ± 0.28 OUR FIT</b>	

 $\Lambda_c(2595)^+ - \Lambda_c^+$  MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>305.79 ± 0.24 OUR FIT</b>				
<b>305.79 ± 0.14 ± 0.20</b>	3.5k	AALTONEN	11H CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
305.6 ± 0.3		<sup>1</sup> BLECHMAN	03	Threshold shift
309.7 ± 0.9 ± 0.4	19	ALBRECHT	97 ARG	$e^+ e^- \approx 10$ GeV
309.2 ± 0.7 ± 0.3	14 ± 4.5	FRABETTI	96 E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
307.5 ± 0.4 ± 1.0	112 ± 17	EDWARDS	95 CLE2	$e^+ e^- \approx 10.5$ GeV

<sup>1</sup> BLECHMAN 03 finds that a more sophisticated treatment than a simple Breit-Wigner for the proximity of the threshold of the dominant decay,  $\Sigma_c(2455)\pi$ , lowers the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass difference by 2 or 3 MeV. The analysis of AALTONEN 11H bears this out.

 $\Lambda_c(2595)^+$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.59 ± 0.30 ± 0.47</b>	3.5k	<sup>2</sup> AALTONEN	11H CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2.9 $\begin{smallmatrix} +2.9 \\ -2.1 \end{smallmatrix}$ $\begin{smallmatrix} +1.8 \\ -1.4 \end{smallmatrix}$	19	ALBRECHT	97 ARG	$e^+ e^- \approx 10$ GeV
3.9 $\begin{smallmatrix} +1.4 \\ -1.2 \end{smallmatrix}$ $\begin{smallmatrix} +2.0 \\ -1.0 \end{smallmatrix}$	112 ± 17	EDWARDS	95 CLE2	$e^+ e^- \approx 10.5$ GeV

<sup>2</sup> AALTONEN 11H treats the three charged modes  $\Lambda_c(2595)^+ \rightarrow \Sigma_c(2455)^+ \pi^+ \pi^-$ ,  $\Sigma_c(2455)^0 \pi^0$ ,  $\Sigma_c(2455)^0 \pi^+$  separately in terms of a common coupling constant  $h_2$  and obtains  $h_2^2 = 0.36 \pm 0.08$ . From this the width is determined.

 $\Lambda_c(2595)^+$  DECAY MODES

$\Lambda_c^+ \pi \pi$  and its submode  $\Sigma_c(2455)\pi$  — the latter just barely — are the only strong decays allowed to an excited  $\Lambda_c^+$  having this mass; and the submode seems to dominate.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda_c^+ \pi^+ \pi^-$	[a] —
$\Gamma_2 \Sigma_c(2455)^{++} \pi^-$	24 ± 7 %
$\Gamma_3 \Sigma_c(2455)^0 \pi^+$	24 ± 7 %
$\Gamma_4 \Lambda_c^+ \pi^+ \pi^- 3\text{-body}$	18 ± 10 %
$\Gamma_5 \Lambda_c^+ \pi^0$	[b] not seen
$\Gamma_6 \Lambda_c^+ \gamma$	not seen

[a] See AALTONEN 11H, Fig. 8, for the calculated ratio of  $\Lambda_c^+ \pi^0 \pi^0$  and  $\Lambda_c^+ \pi^+ \pi^-$  partial widths as a function of the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass difference. At our value of the mass difference, the ratio is about 4.

[b] A test that the isospin is indeed 0, so that the particle is indeed a  $\Lambda_c^+$ .

 $\Lambda_c(2595)^+$  BRANCHING RATIOS

$\Gamma(\Sigma_c(2455)^{++}\pi^-)/\Gamma(\Lambda_c^+\pi^+\pi^-)$				$\Gamma_2/\Gamma_1$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.36±0.10 OUR AVERAGE</b>				
0.37±0.12±0.13	ALBRECHT	97 ARG	$e^+e^- \approx 10$ GeV	
0.36±0.09±0.09	EDWARDS	95 CLE2	$e^+e^- \approx 10.5$ GeV	

$\Gamma(\Sigma_c(2455)^0 \pi^+)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$	$\Gamma_3/\Gamma_1$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.37±0.10 OUR AVERAGE</b>			
0.29±0.10±0.11	ALBRECHT	97 ARG	$e^+e^- \approx 10$ GeV
0.42±0.09±0.09	EDWARDS	95 CLE2	$e^+e^- \approx 10.5$ GeV

$\frac{[\Gamma(\Sigma_c(2455)^{++}\pi^-) + \Gamma(\Sigma_c(2455)^0\pi^+)]/\Gamma(\Lambda_c^+\pi^+\pi^-)}{(\Gamma_2 + \Gamma_3)/\Gamma_1}$				
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$0.66^{+0.13}_{-0.16} \pm 0.07$		ALBRECHT	97	ARG	$e^+e^- \approx 10$ GeV
$>0.51$	90	<sup>3</sup> FRABETTI	96	E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV

<sup>3</sup> The results of FRABETTI 96 are consistent with this ratio being 100%.

$\Gamma(\Lambda_c^+ \pi^0)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_5/\Gamma_1$
$\Lambda_c^+ \pi^0$ decay is forbidden by isospin conservation if this state is in fact a $\Lambda_c$ .					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;3.53</b>	90	EDWARDS	95 CLE2	$e^+ e^- \approx 10.5$ GeV	

$\Gamma(\Lambda_c^+ \gamma)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$				$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.98	90	EDWARDS	95 CLE2	$e^+ e^- \approx 10.5$ GeV

 $\Lambda_c(2595)^+$  REFERENCES

AALTONEN	11H	PR D84 012003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
BLECHMAN	03	PR D67 074033	A.E. Blechman <i>et al.</i>	(JHU, FLOR)
ALBRECHT	97	PL B402 207	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
FRABETTI	96	PL B365 461	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
EDWARDS	95	PRL 74 3331	K.W. Edwards <i>et al.</i>	(CLEO Collab.)

$$\Lambda_c(2625)^+$$

$$J(P) = 0(\frac{3}{2}^-) \text{ Status: } ***$$

The spin-parity has not been measured but is expected to be  $3/2^-$ : this is presumably the charm counterpart of the strange  $\Lambda(1520)$ .

 $\Lambda_c(2625)^+$  MASS

The mass is obtained from the  $\Lambda_c(2625)^+ - \Lambda_c^+$  mass-difference measurements below.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2628.11 ± 0.19 OUR FIT</b>				Error includes scale factor of 1.1.
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2626.6 ± 0.5 ± 1.5	42 ± 9	ALBRECHT	93F ARG	See ALBRECHT 97

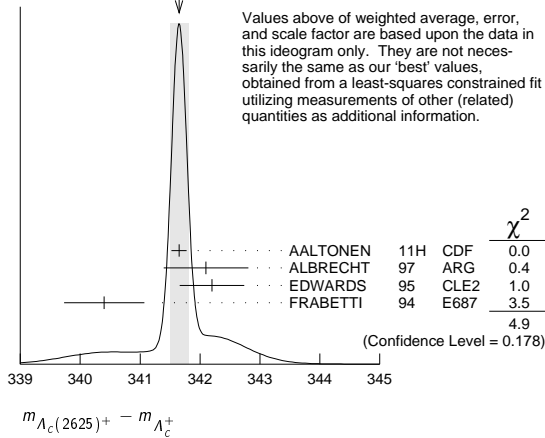
# Baryon Particle Listings

## $\Lambda_c(2625)^+, \Lambda_c(2765)^+, \Lambda_c(2880)^+$

### $\Lambda_c(2625)^+ - \Lambda_c^+$ MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>341.65 \pm 0.13</math> OUR FIT</b>	Error includes scale factor of 1.1.			
<b><math>341.65 \pm 0.15</math> OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.			
$341.65 \pm 0.04 \pm 0.12$	6.2k	AALTONEN	11H	CDF $p\bar{p}$ at 1.96 TeV
$342.1 \pm 0.5 \pm 0.5$	51	ALBRECHT	97	ARG $e^+e^- \approx 10$ GeV
$342.2 \pm 0.2 \pm 0.5$	$245 \pm 19$	EDWARDS	95	CLE2 $e^+e^- \approx 10.5$ GeV
$340.4 \pm 0.6 \pm 0.3$	$40 \pm 9$	FRABETTI	94	E687 $\gamma$ Be, $\bar{E}_\gamma = 220$ GeV

WEIGHTED AVERAGE  
 $341.65 \pm 0.15$  (Error scaled by 1.3)



### $\Lambda_c(2625)^+$ WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.97</b>	90	6.2k	AALTONEN	11H	CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<1.9	90	$245 \pm 19$	EDWARDS	95	CLE2 $e^+e^- \approx 10.5$ GeV
<3.2	90		ALBRECHT	93F	ARG $e^+e^- \approx \mathcal{T}(4S)$

### $\Lambda_c(2625)^+$ DECAY MODES

$\Lambda_c^+ \pi \pi$  and its submode  $\Sigma(2455)\pi$  are the only strong decays allowed to an excited  $\Lambda_c^+$  having this mass.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1$ $\Lambda_c^+ \pi^+ \pi^-$	[a] $\approx 67\%$	
$\Gamma_2$ $\Sigma_c(2455)^{++} \pi^-$	<5	90%
$\Gamma_3$ $\Sigma_c(2455)^0 \pi^+$	<5	90%
$\Gamma_4$ $\Lambda_c^+ \pi^+ \pi^-$ 3-body	large	
$\Gamma_5$ $\Lambda_c^+ \pi^0$	[b] not seen	
$\Gamma_6$ $\Lambda_c^+ \gamma$	not seen	

[a] See AALTONEN 11H, Fig. 8, for the calculated ratio of  $\Lambda_c^+ \pi^0 \pi^0$  and  $\Lambda_c^+ \pi^+ \pi^-$  partial widths as a function of the  $\Lambda_c(2595)^+ - \Lambda_c^+$  mass difference. At our value of the mass difference, the ratio is about 4.

[b] A test that the isospin is indeed 0, so that the particle is indeed a  $\Lambda_c^+$ .

### $\Lambda_c(2625)^+$ BRANCHING RATIOS

$\Gamma(\Sigma_c(2455)^{++} \pi^-)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_2/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.08</b>	90	EDWARDS	95	CLE2 $e^+e^- \approx 10.5$ GeV	

$\Gamma(\Sigma_c(2455)^0 \pi^+)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_3/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.07</b>	90	EDWARDS	95	CLE2 $e^+e^- \approx 10.5$ GeV	

$[\Gamma(\Sigma_c(2455)^{++} \pi^-) + \Gamma(\Sigma_c(2455)^0 \pi^+)]/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$(\Gamma_2 + \Gamma_3)/\Gamma_1$
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.36	90		FRABETTI	94	E687 $\gamma$ Be, $\bar{E}_\gamma = 220$ GeV
$0.46 \pm 0.14$	21		ALBRECHT	93F	ARG $e^+e^- \approx \mathcal{T}(4S)$

$\Gamma(\Lambda_c^+ \pi^+ \pi^- \text{ 3-body})/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_4/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$0.54 \pm 0.14$	16	ALBRECHT	93F	ARG $e^+e^- \approx \mathcal{T}(4S)$	

$\Gamma(\Lambda_c^+ \pi^0)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_5/\Gamma_1$
$\Lambda_c^+ \pi^0$ decay is forbidden by isospin conservation if this state is in fact a $\Lambda_c$ .					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.91</b>	90	EDWARDS	95	CLE2 $e^+e^- \approx 10.5$ GeV	

$\Gamma(\Lambda_c^+ \gamma)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$					$\Gamma_6/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.52</b>	90	EDWARDS	95	CLE2 $e^+e^- \approx 10.5$ GeV	

### $\Lambda_c(2625)^+$ REFERENCES

AALTONEN	11H	PR D84 012003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ALBRECHT	97	PL B402 207	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
EDWARDS	95	PRL 74 3331	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
FRABETTI	94	PRL 72 961	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALBRECHT	93F	PL B317 227	H. Albrecht <i>et al.</i>	(ARGUS Collab.)

$\Lambda_c(2765)^+$   
or  $\Sigma_c(2765)$

$I(J^P) = ?(??)$  Status: \*

OMITTED FROM SUMMARY TABLE

A broad, statistically significant peak ( $997^{+141}_{-129}$  events) seen in  $\Lambda_c^+ \pi^+ \pi^-$ . However, nothing at all is known about its quantum numbers, including whether it is a  $\Lambda_c^+$  or a  $\Sigma_c$ , or whether the width might be due to overlapping states.

### $\Lambda_c(2765)^+$ MASS

The mass is obtained from the  $\Lambda_c(2765)^+ - \Lambda_c^+$  mass-difference measurement below.

VALUE (MeV)	DOCUMENT ID
<b><math>2766.6 \pm 2.4</math> OUR FIT</b>	

### $\Lambda_c(2765)^+ - \Lambda_c^+$ MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>480.1 \pm 2.4</math> OUR FIT</b>				
<b><math>480.1 \pm 2.4</math></b>	$997^{+141}_{-129}$	ARTUSO	01	CLE2 $e^+e^- \approx \mathcal{T}(4S)$

### $\Lambda_c(2765)^+$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>50</b>	ARTUSO	01	CLE2 $e^+e^- \approx \mathcal{T}(4S)$

### $\Lambda_c(2765)^+$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\Lambda_c^+ \pi^+ \pi^-$	seen

### $\Lambda_c(2765)^+$ REFERENCES

ARTUSO	01	PRL 86 4479	M. Artuso <i>et al.</i>	(CLEO Collab.)
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$\Lambda_c(2880)^+$

$I(J^P) = 0(\frac{5}{2}^+)$  Status: \*\*\*

A narrow peak seen in  $\Lambda_c^+ \pi^+ \pi^-$  and in  $pD^0$ . It is not seen in  $pD^+$ , and therefore it is probably a  $\Lambda_c^+$  and not a  $\Sigma_c$ . The evidence for spin 5/2 comes from the  $\Sigma_c(2455)\pi$  decay angular distribution, and the evidence for parity + comes from agreement of the  $\Sigma_c(2520)/\Sigma_c(2455)$  branching ratio with a prediction of heavy quark symmetry (see MIZUK 07).

### $\Lambda_c(2880)^+$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2881.53 \pm 0.35</math> OUR FIT</b>				
<b><math>2881.50 \pm 0.35</math> OUR AVERAGE</b>				
$2881.9 \pm 0.1 \pm 0.5$	$2.8k \pm 190$	AUBERT	07	BABR in $pD^0$
$2881.2 \pm 0.2 \pm 0.4$	$690 \pm 50$	MIZUK	07	BELL in $\Sigma_c(2455)^0, ++ \pi^\pm$

See key on page 601

Baryon Particle Listings

$\Lambda_c(2880)^+$ ,  $\Lambda_c(2940)^+$ ,  $\Sigma_c(2455)$

$\Lambda_c(2880)^+ - \Lambda_c^+$ MASS DIFFERENCE				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>595.1±0.4 OUR FIT</b>				
<b>596 ±1 ±2</b>	350 <sup>+57</sup> <sub>-55</sub>	ARTUSO	01	CLE2 in $\Lambda_c^+ \pi^+ \pi^-$

$\Lambda_c(2880)^+$ WIDTH					
VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.8±1.1 OUR AVERAGE</b>					
5.8±1.5±1.1		2.8k±190	AUBERT	07	BABR in $pD^0$
5.8±0.7±1.1		690 ± 50	MIZUK	07	BELL in $\Sigma_c(2455)^{0,++} \pi^\pm$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<8	90		ARTUSO	01	CLEO in $\Lambda_c^+ \pi^+ \pi^-$

$\Lambda_c(2880)^+$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\Lambda_c^+ \pi^+ \pi^-$	seen
$\Gamma_2$ $\Sigma_c(2455)^{0,++} \pi^\pm$	seen
$\Gamma_3$ $\Sigma_c(2520)^{0,++} \pi^\pm$	seen
$\Gamma_4$ $pD^0$	seen

$\Lambda_c(2880)^+$ BRANCHING RATIOS				
$\Gamma(\Sigma_c(2455)^{0,++} \pi^\pm)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$				$\Gamma_2/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.392±0.031 OUR AVERAGE</b>				Error includes scale factor of 1.3.
0.404±0.021±0.014		MIZUK	07	BELL in $\Sigma_c(2455)^{0,++} \pi^\pm$
0.31 ±0.06 ±0.03	96	ARTUSO	01	CLE2 $e^+ e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.11	90	ARTUSO	01	CLE2 $e^+ e^- \approx \gamma(4S)$
$\Gamma(\Sigma_c(2520)^{0,++} \pi^\pm)/\Gamma(\Lambda_c^+ \pi^+ \pi^-)$				$\Gamma_3/\Gamma_1$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.091±0.025±0.010</b>		MIZUK	07	BELL in $\Sigma_c(2455)^{0,++} \pi^\pm$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.11	90	ARTUSO	01	CLE2 $e^+ e^- \approx \gamma(4S)$
$\Gamma(\Sigma_c(2520)^{0,++} \pi^\pm)/\Gamma(\Sigma_c(2455)^{0,++} \pi^\pm)$				$\Gamma_3/\Gamma_2$
VALUE	DOCUMENT ID	TECN	COMMENT	
0.225±0.062±0.025	<sup>1</sup> MIZUK	07	BELL in $\Sigma_c(2455)^{0,++} \pi^\pm$	
<sup>1</sup> This MIZUK 07 ratio is redundant with MIZUK 07 ratios given above.				

$\Lambda_c(2880)^+$ REFERENCES				
AUBERT	07	PRL 98 012001	B. Aubert <i>et al.</i>	(BABAR Collab.)
MIZUK	07	PRL 98 262001	R. Mizuk <i>et al.</i>	(BELLE Collab.)
ARTUSO	01	PRL 86 4479	M. Artuso <i>et al.</i>	(CLEO Collab.)

<div><math>\Lambda_c(2940)^+</math></div>	$I(J^P) = 0(?^?)$ Status: ***
A fairly narrow peak of good statistical significance first seen in the $pD^0$ mass spectrum. It is not seen in $pD^+$ , and thus it is probably a $\Lambda_c^+$ and not a $\Sigma_c$ . It is also seen in $\Sigma_c(2455)^{0,++} \pi^\pm$ .	

$\Lambda_c(2940)^+$ MASS				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2939.3<sup>+1.4</sup><sub>-1.5</sub> OUR AVERAGE</b>				
2939.8±1.3±1.0	2280 ± 310	AUBERT	07	BABR in $pD^0$
2938.0±1.3 <sup>+2.0</sup> <sub>-4.0</sub>	220 <sup>+80</sup> <sub>-60</sub>	MIZUK	07	BELL in $\Sigma_c(2455)^{0,++} \pi^\pm$

$\Lambda_c(2940)^+$ WIDTH				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17 <sup>+8</sup><sub>-6</sub> OUR AVERAGE</b>				
17.5±5.2± 5.9	2280 ± 310	AUBERT	07	BABR in $pD^0$
13 <sup>+8</sup> <sub>-5</sub> <sup>+27</sup> <sub>-7</sub>	220 <sup>+80</sup> <sub>-60</sub>	MIZUK	07	BELL in $\Sigma_c(2455)^{0,++} \pi^\pm$

$\Lambda_c(2940)^+$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $pD^0$	seen
$\Gamma_2$ $\Sigma_c(2455)^{0,++} \pi^\pm$	seen

$\Lambda_c(2940)^+$ REFERENCES				
AUBERT	07	PRL 98 012001	B. Aubert <i>et al.</i>	(BABAR Collab.)
MIZUK	07	PRL 98 262001	R. Mizuk <i>et al.</i>	(BELLE Collab.)

<div><math>\Sigma_c(2455)</math></div>	$I(J^P) = 1(\frac{1}{2}^+)$ Status: ***
The angular distribution of $B^- \rightarrow \Sigma_c(2455)^0 \bar{p}$ favors $J = 1/2$ (as the quark model predicts). $J = 3/2$ is excluded by more than four $\sigma$ see AUBERT 08BN.	

$\Sigma_c(2455)$ MASSES	
The masses are obtained from the mass-difference measurements that follow.	

$\Sigma_c(2455)^{++}$ MASS	
VALUE (MeV)	DOCUMENT ID
<b>2453.97±0.14 OUR FIT</b>	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$\Sigma_c(2455)^+$ MASS	
VALUE (MeV)	DOCUMENT ID
<b>2452.9±0.4 OUR FIT</b>	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$\Sigma_c(2455)^0$ MASS	
VALUE (MeV)	DOCUMENT ID
<b>2453.75±0.14 OUR FIT</b>	
• • • We do not use the following data for averages, fits, limits, etc. • • •	

$\Sigma_c(2455) - \Lambda_c^+$ MASS DIFFERENCES				
$m_{\Sigma_c^{++}} - m_{\Lambda_c^+}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>167.510± 0.017 OUR FIT</b>				
<b>167.510± 0.022 OUR AVERAGE</b>				
167.51 ± 0.01 ±0.02	36k	LEE	14	BELL $e^+ e^-$ at $\gamma(4S)$
167.44 ± 0.04 ±0.12	13.8k	AALTONEN	11H	CDF $p\bar{p}$ at 1.96 TeV
167.4 ± 0.1 ±0.2	2k	ARTUSO	02	CLE2 $e^+ e^- \approx \gamma(4S)$
167.35 ± 0.19 ±0.12	461	LINK	00c	FOCS $\gamma A, \bar{E}_\gamma$ 180 GeV
167.76 ± 0.29 ±0.15	122	AITALA	96B	E791 $\pi^- N$ , 500 GeV
167.6 ± 0.6 ±0.6	56	FRABETTI	96	E687 $\gamma\text{Be}, \bar{E}_\gamma \approx 220$ GeV
168.2 ± 0.3 ±0.2	126	CRAWFORD	93	CLE2 $e^+ e^- \approx \gamma(4S)$
167.8 ± 0.4 ±0.3	54	BOWCOCK	89	CLEO $e^+ e^-$ 10 GeV
168.2 ± 0.5 ±1.6	92	ALBRECHT	88D	ARG $e^+ e^-$ 10 GeV
167.4 ± 0.5 ±2.0	46	DIESBURG	87	SPEC $nA \sim 600$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
167 ± 1	2	JONES	87	HBC $\nu p$ in BEBC
166 ± 1	1	BOSETTI	82	HBC See JONES 87
168 ± 3	6	BALTAY	79	HLBC $\nu$ Ne-H in 15-ft
166 ±15	1	CAZZOLI	75	HBC $\nu p$ in BNL 7-ft

$m_{\Sigma_c^+} - m_{\Lambda_c^+}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>166.4±0.4 OUR FIT</b>				
<b>166.4±0.2±0.3</b>	661	AMMAR	01	CLE2 $e^+ e^- \approx \gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
168.5±0.4±0.2	111	CRAWFORD	93	CLE2 See AMMAR 01
168 ±3	1	CALICCHIO	80	HBC $\nu p$ in BEBC-TST

$m_{\Sigma_c^0} - m_{\Lambda_c^+}$				
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>167.290±0.017 OUR FIT</b>				
<b>167.290±0.022 OUR AVERAGE</b>				
167.29 ±0.01 ±0.02	32k	LEE	14	BELL $e^+ e^-$ at $\gamma(4S)$
167.28 ±0.03 ±0.12	15.9k	AALTONEN	11H	CDF $p\bar{p}$ at 1.96 TeV
167.2 ±0.1 ±0.2	2k	ARTUSO	02	CLE2 $e^+ e^- \approx \gamma(4S)$
167.38 ±0.21 ±0.13	362	LINK	00c	FOCS $\gamma A, \bar{E}_\gamma$ 180 GeV
167.38 ±0.29 ±0.15	143	AITALA	96B	E791 $\pi^- N$ , 500 GeV
167.8 ±0.6 ±0.2		ALEEV	96	SPEC $n$ nucleus, 50 GeV/c
166.6 ±0.5 ±0.6	69	FRABETTI	96	E687 $\gamma\text{Be}, \bar{E}_\gamma \approx 220$ GeV
167.1 ±0.3 ±0.2	124	CRAWFORD	93	CLE2 $e^+ e^- \approx \gamma(4S)$
168.4 ±1.0 ±0.3	14	ANJOS	89D	E691 $\gamma\text{Be}$ 90–260 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
167.9 ±0.5 ±0.3	48	<sup>1</sup> BOWCOCK	89	CLEO $e^+ e^-$ 10 GeV
167.0 ±0.5 ±1.6	70	<sup>1</sup> ALBRECHT	88D	ARG $e^+ e^-$ 10 GeV
178.2 ±0.4 ±2.0	85	<sup>2</sup> DIESBURG	87	SPEC $nA \sim 600$ GeV
163 ±2	1	AMMAR	86	EMUL $\nu A$

<sup>1</sup> This result enters the fit through  $m_{\Sigma_c^{++}} - m_{\Sigma_c^-}$  given below.

<sup>2</sup> See the note on DIESBURG 87 in the  $m_{\Sigma_c^{++}} - m_{\Sigma_c^0}$  section below.

Baryon Particle Listings

$\Sigma_c(2455), \Sigma_c(2520)$

$\Sigma_c(2455)$  MASS DIFFERENCES

$m_{\Sigma_c^{++}} - m_{\Sigma_c^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>0.220±0.013 OUR FIT</b>			
<b>0.221±0.014 OUR AVERAGE</b>			
0.22 ± 0.01 ± 0.01	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
0.2 ± 0.1 ± 0.1	ARTUSO	02	CLE2 $e^+e^- \approx \Upsilon(4S)$
− 0.03 ± 0.28 ± 0.11	LINK	00c	FOCS $\gamma A, \overline{E}_\gamma$ 180 GeV
0.38 ± 0.40 ± 0.15	AITALA	96b	E791 $\pi^- N$ , 500 GeV
1.1 ± 0.4 ± 0.1	CRAWFORD	93	CLE2 $e^+e^- \approx \Upsilon(4S)$
− 0.1 ± 0.6 ± 0.1	BOWCOCK	89	CLEO $e^+e^-$ 10 GeV
1.2 ± 0.7 ± 0.3	ALBRECHT	88d	ARG $e^+e^- \sim$ 10 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
−10.8 ± 2.9	<sup>3</sup> DIESBURG	87	SPEC $nA \sim$ 600 GeV
<sup>3</sup> DIESBURG 87 is completely incompatible with the other experiments, which is surprising since it agrees with them about $m_{\Sigma_c(2455)^{++}} - m_{\Lambda_c^+}$ . We go with the majority here.			

$m_{\Sigma_c^+} - m_{\Sigma_c^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>−0.9±0.4 OUR FIT</b>			
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.4 ± 0.5 ± 0.3	CRAWFORD	93	CLE2 See AMMAR 01

$\Sigma_c(2455)$  WIDTHS

$\Sigma_c(2455)^{++}$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.89±0.09 OUR AVERAGE</b>				Error includes scale factor of 1.1.
1.84 ± 0.04 ± 0.07 − 0.20	36k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
2.34 ± 0.13 ± 0.45	13.8k	AALTONEN	11H	CDF $p\overline{p}$ at 1.96 TeV
2.3 ± 0.2 ± 0.3	2k	ARTUSO	02	CLE2 $e^+e^- \approx \Upsilon(4S)$
2.05 ± 0.41 ± 0.38 − 0.38	1110	LINK	02	FOCS $\gamma A, \overline{E}_\gamma \approx$ 180 GeV

$\Sigma_c(2455)^+$  WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.6</b>	90	661	AMMAR	01	CLE2 $e^+e^- \approx \Upsilon(4S)$

$\Sigma_c(2455)^0$  WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.83±0.11 OUR AVERAGE</b>				Error includes scale factor of 1.2.
1.76 ± 0.04 ± 0.09 − 0.21	32k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
1.65 ± 0.11 ± 0.49	15.9k	AALTONEN	11H	CDF $p\overline{p}$ at 1.96 TeV
2.6 ± 0.5 ± 0.3		AUBERT	08BN	BABR $B^- \rightarrow \overline{p}\Lambda_c^+\pi^-$
2.5 ± 0.2 ± 0.3	2k	ARTUSO	02	CLE2 $e^+e^- \approx \Upsilon(4S)$
1.55 ± 0.41 ± 0.38 − 0.37	913	LINK	02	FOCS $\gamma A, \overline{E}_\gamma \approx$ 180 GeV

$\Sigma_c(2455)$  DECAY MODES

$\Lambda_c^+\pi$  is the only strong decay allowed to a  $\Sigma_c$  having this mass.

Mode	Fraction ( $\Gamma_j/\Gamma$ )
$\Gamma_1 \quad \Lambda_c^+\pi$	$\approx$ 100 %

$\Sigma_c(2455)$  REFERENCES

LEE	14	PR D89 091102	S.-H. Lee <i>et al.</i>	(BELLE Collab.)
AALTONEN	11H	PR D84 012003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AUBERT	08BN	PR D78 112003	B. Aubert <i>et al.</i>	(BABAR Collab.)
ARTUSO	02	PR D65 071101	M. Artuso <i>et al.</i>	(CLEO Collab.)
LINK	02	PL B525 205	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
AMMAR	01	PRL 86 1167	R. Ammar <i>et al.</i>	(CLEO Collab.)
LINK	00c	PL B488 218	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
AITALA	96b	PL B379 292	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
ALEEV	96	JINRR C 3-77 31	A.N. Aleev <i>et al.</i>	(Serpukhov EXCHARM Collab.)
FRABETTI	96	PL B365 461	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
CRAWFORD	93	PRL 71 3259	G. Crawford <i>et al.</i>	(CLEO Collab.)
ANJOS	89D	PRL 62 1721	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BOWCOCK	89	PRL 62 1240	T.J.V. Bowcock <i>et al.</i>	(CLEO Collab.)
ALBRECHT	88D	PL B211 489	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DIESBURG	87	PRL 59 2711	M. Diesburg <i>et al.</i>	(FNAL E400 Collab.)
JONES	87	ZPHY C36 593	G.T. Jones <i>et al.</i>	(CERN WA21 Collab.)
AMMAR	86	JETPL 43 515	R. Ammar <i>et al.</i>	(ITEP)
		Translated from ZETFP 43 401		
BOSETTI	82	PL 109B 234	P.C. Bosetti <i>et al.</i>	(AACH3, BONN, CERN+)
CALICCHIO	80	PL 93B 521	M. Calicchio <i>et al.</i>	(BARI, BIRM, BRUX+)
BALTAY	79	PRL 42 1721	C. Baltay <i>et al.</i>	(COLU, BNL)
CAZZOLI	75	PRL 34 1125	E.G. Cazzoli <i>et al.</i>	(BNL)

$\Sigma_c(2520)$

$I(J^P) = 1(\frac{3}{2}^+)$  Status: \*\*\*

Seen in the  $\Lambda_c^+\pi^\pm$  mass spectrum. The natural assignment is that this is the  $J^P = 3/2^+$  excitation of the  $\Sigma_c(2455)$ , the charm counterpart of the  $\Sigma(1385)$ , but neither  $J$  nor  $P$  has been measured.

$\Sigma_c(2520)$  MASSES

The masses are obtained from the mass-difference measurements that follow.

$\Sigma_c(2520)^{++}$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2518.41±0.21 OUR FIT</b>				Error includes scale factor of 1.1.
2530 ± 5 ± 5	6	<sup>1</sup> AMMOSOV	93	HLBC $\nu p \rightarrow \mu^- \Sigma_c(2530)^{++}$
<sup>1</sup> AMMOSOV 93 sees a cluster of 6 events and estimates the background to be 1 event.				

$\Sigma_c(2520)^+$  MASS

VALUE (MeV)	DOCUMENT ID
<b>2517.5±2.3 OUR FIT</b>	

$\Sigma_c(2520)^0$  MASS

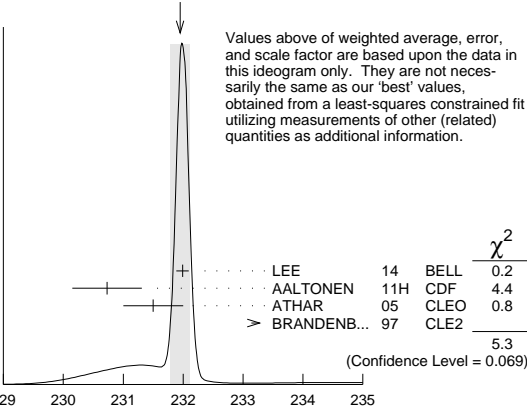
VALUE (MeV)	DOCUMENT ID
<b>2518.48±0.20 OUR FIT</b>	

$\Sigma_c(2520)$  MASS DIFFERENCES

$m_{\Sigma_c(2520)^{++}} - m_{\Lambda_c^+}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>231.95±0.17 OUR FIT</b>				Error includes scale factor of 1.3.
<b>231.95±0.16 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
231.99 ± 0.10 ± 0.02	44k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
230.73 ± 0.56 ± 0.16	8.8k	AALTONEN	11H	CDF $p\overline{p}$ at 1.96 TeV
231.5 ± 0.4 ± 0.3	1.3k	ATHAR	05	CLEO $e^+e^-$ , 9.4–11.5 GeV
234.5 ± 1.1 ± 0.8	677	BRANDENB...	97	CLE2 $e^+e^- \approx \Upsilon(4S)$

WEIGHTED AVERAGE  
231.95±0.16 (Error scaled by 1.6)



$m_{\Sigma_c(2520)^{++}} - m_{\Lambda_c^+}$  (MeV)

$m_{\Sigma_c(2520)^+} - m_{\Lambda_c^+}$

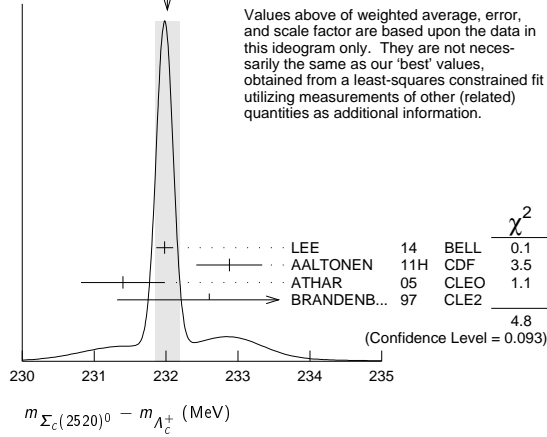
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>231.0±2.3 OUR FIT</b>				
<b>231.0±1.1±2.0</b>	327	AMMAR	01	CLE2 $e^+e^- \approx \Upsilon(4S)$

$m_{\Sigma_c(2520)^0} - m_{\Lambda_c^+}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>232.02±0.15 OUR FIT</b>				Error includes scale factor of 1.3.
<b>232.02±0.17 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
231.98 ± 0.11 ± 0.04	41k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
232.88 ± 0.43 ± 0.16	9.0k	AALTONEN	11H	CDF $p\overline{p}$ at 1.96 TeV
231.4 ± 0.5 ± 0.3	1.3k	ATHAR	05	CLEO $e^+e^-$ , 9.4–11.5 GeV
232.6 ± 1.0 ± 0.8	504	BRANDENB...	97	CLE2 $e^+e^- \approx \Upsilon(4S)$

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## Baryon Particle Listings

 $\Sigma_c(2520)$ ,  $\Sigma_c(2800)$ WEIGHTED AVERAGE  
232.02±0.17 (Error scaled by 1.5) $m_{\Sigma_c(2520)^{++}} - m_{\Sigma_c(2520)^0}$ 

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>0.01±0.15±0.03</b>	44/41k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.1 ± 0.8 ± 0.3		<sup>2</sup> ATHAR	05	CLEO $e^+e^-$ , 9.4–11.5 GeV
1.9 ± 1.4 ± 1.0		<sup>3</sup> BRANDENB...	97	CLE2 $e^+e^- \approx \Upsilon(4S)$

<sup>2</sup> This ATHAR 05 result is redundant with measurements in earlier entries.<sup>3</sup> This BRANDENBURG 97 result is redundant with measurements in earlier entries. $\Sigma_c(2520)$  WIDTHS $\Sigma_c(2520)^{++}$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>14.78<sup>+0.30</sup><sub>-0.40</sub></b>	<b>OUR AVERAGE</b>			
14.77±0.25 <sup>+0.18</sup> <sub>-0.30</sub>	44k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
15.03±2.12±1.36	8.8k	AALTONEN	11H	CDF $p\bar{p}$ at 1.96 TeV
14.4 <sup>+1.6</sup> <sub>-1.5</sub> ± 1.4	1.3k	ATHAR	05	CLEO $e^+e^-$ , 9.4–11.5 GeV
17.9 <sup>+3.8</sup> <sub>-3.2</sub> ± 4.0	677	BRANDENB...	97	CLE2 $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2520)^+$  WIDTH

VALUE (MeV)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<b>&lt;17</b>	90	327	AMMAR	01	CLE2 $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2520)^0$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>15.3<sup>+0.4</sup><sub>-0.5</sub></b>	<b>OUR AVERAGE</b>			
15.41±0.41 <sup>+0.20</sup> <sub>-0.32</sub>	41k	LEE	14	BELL $e^+e^-$ at $\Upsilon(4S)$
12.51±1.82±1.37	9.0k	AALTONEN	11H	CDF $p\bar{p}$ at 1.96 TeV
16.6 <sup>+1.9</sup> <sub>-1.7</sub> ± 1.4	1.3k	ATHAR	05	CLEO $e^+e^-$ , 9.4–11.5 GeV
13.0 <sup>+3.7</sup> <sub>-3.0</sub> ± 4.0	504	BRANDENB...	97	CLE2 $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2520)$  DECAY MODES $\Lambda_c^+ \pi$  is the only strong decay allowed to a  $\Sigma_c$  having this mass.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda_c^+ \pi$	≈ 100 %

 $\Sigma_c(2520)$  REFERENCES

LEE	14	PR D89 091102	S.-H. Lee <i>et al.</i>	(BELLE Collab.)
AALTONEN	11H	PR D84 012003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ATHAR	05	PR D71 051101	S.B. Athar <i>et al.</i>	(CLEO Collab.)
AMMAR	01	PRL 86 1167	R. Ammar <i>et al.</i>	(CLEO Collab.)
BRANDENB...	97	PRL 78 2304	G. Brandenburg <i>et al.</i>	(CLEO Collab.)
AMMOS OV	93	JETPL 58 247	V.V. Ammosov <i>et al.</i>	(SERP)
Translated from ZETFP 58 241.				

 $\Sigma_c(2800)$  $I(J^P) = 1(?)^?$  Status: \*\*\*Seen in the  $\Lambda_c^+ \pi^+$ ,  $\Lambda_c^+ \pi^0$ , and  $\Lambda_c^+ \pi^-$  mass spectra. $\Sigma_c(2800)$  MASSES

The charged ++ and + masses are obtained from the mass-difference measurements that follow. The neutral mass is dominated by the mass-difference measurement, but is pulled up somewhat by the less well-determined but considerably higher direct-mass measurement. It is possible, in fact, that AUBERT 08BN is seeing a different  $\Sigma_c$ .

 $\Sigma_c(2800)^{++}$  MASS

VALUE (MeV)	DOCUMENT ID
<b>2801<sup>+4</sup><sub>-6</sub></b>	<b>OUR FIT</b>

 $\Sigma_c(2800)^+$  MASS

VALUE (MeV)	DOCUMENT ID
<b>2792<sup>+14</sup><sub>-5</sub></b>	<b>OUR FIT</b>

 $\Sigma_c(2800)^0$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2806<sup>+5</sup><sub>-7</sub></b>	<b>OUR FIT</b>		Error includes scale factor of 1.3.
<b>2846±8±10</b>	AUBERT	08BN BABR	$B^- \rightarrow \bar{p} \Lambda_c^+ \pi^-$

 $\Sigma_c(2800)$  MASS DIFFERENCES $m_{\Sigma_c(2800)^{++}} - m_{\Lambda_c^+}$ 

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>514<sup>+4</sup><sub>-6</sub></b>	<b>OUR FIT</b>			
<b>514.5<sup>+3.4</sup><sub>-3.1</sub></b>	2810 <sup>+1090</sup> <sub>-775</sub>	MIZUK	05	BELL $e^+e^- \approx \Upsilon(4S)$

 $m_{\Sigma_c(2800)^+} - m_{\Lambda_c^+}$ 

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>505<sup>+14</sup><sub>-5</sub></b>	<b>OUR FIT</b>			
<b>505.4<sup>+5.8</sup><sub>-4.6</sub></b>	1540 <sup>+1750</sup> <sub>-1050</sub>	MIZUK	05	BELL $e^+e^- \approx \Upsilon(4S)$

 $m_{\Sigma_c(2800)^0} - m_{\Lambda_c^+}$ 

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>519<sup>+5</sup><sub>-7</sub></b>	<b>OUR FIT</b>			
<b>515.4<sup>+3.2</sup><sub>-3.1</sub></b>	2240 <sup>+1300</sup> <sub>-740</sub>	MIZUK	05	BELL $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2800)$  WIDTHS $\Sigma_c(2800)^{++}$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>75<sup>+18</sup><sub>-13</sub></b>	2810 <sup>+1090</sup> <sub>-775</sub>	MIZUK	05	BELL $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2800)^+$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>62<sup>+37</sup><sub>-23</sub></b>	1540 <sup>+1750</sup> <sub>-1050</sub>	MIZUK	05	BELL $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2800)^0$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b>72<sup>+22</sup><sub>-15</sub></b>	<b>OUR AVERAGE</b>			
86 <sup>+33</sup> <sub>-22</sub> ± 12		AUBERT	08BN BABR	$B^- \rightarrow \bar{p} \Lambda_c^+ \pi^-$
61 <sup>+18</sup> <sub>-13</sub> ± 22	2240 <sup>+1300</sup> <sub>-740</sub>	MIZUK	05	BELL $e^+e^- \approx \Upsilon(4S)$

 $\Sigma_c(2800)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda_c^+ \pi$	seen

 $\Sigma_c(2800)$  REFERENCES

AUBERT	08BN	PR D78 112003	B. Aubert <i>et al.</i>	(BABAR Collab.)
MIZUK	05	PRL 94 122002	R. Mizuk <i>et al.</i>	(BELLE Collab.)

# Baryon Particle Listings



$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \*\*\*

According to the quark model, the  $\Xi_c^+$  (quark content  $usc$ ) and  $\Xi_c^0$  form an isospin doublet, and the spin-parity ought to be  $J^P = 1/2^+$ . None of  $I$ ,  $J$ , or  $P$  has actually been measured.

## $\Xi_c^+$ MASS

The fit uses the  $\Xi_c^+$  and  $\Xi_c^0$  mass and mass-difference measurements.

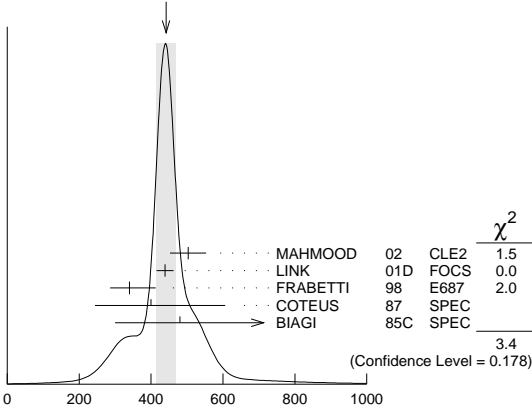
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2467.93<math>\pm</math> 0.28 0.40 OUR FIT</b>				
<b>2467.89<math>\pm</math> 0.34 0.50 OUR AVERAGE</b>				
2468.00 $\pm$ 0.18 $\pm$ 0.51	5.1k	AALTONEN	14B CDF	$p\bar{p}$ at 1.96 TeV
2468.1 $\pm$ 0.4 $\pm$ 0.2 1.4	4.9k	<sup>1</sup> LESIAK	05 BELL	$e^+e^-$ , $\Upsilon(4S)$
2465.8 $\pm$ 1.9 $\pm$ 2.5	90	FRABETTI	98 E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
2467.0 $\pm$ 1.6 $\pm$ 2.0	147	EDWARDS	96 CLE2	$e^+e^- \approx \Upsilon(4S)$
2465.1 $\pm$ 3.6 $\pm$ 1.9	30	ALBRECHT	90f ARG	$e^+e^-$ at $\Upsilon(4S)$
2467 $\pm$ 3 $\pm$ 4	23	ALAM	89 CLEO	$e^+e^-$ 10.6 GeV
2466.5 $\pm$ 2.7 $\pm$ 1.2	5	BARLAG	89c ACCM	$\pi^-$ Cu 230 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2464.4 $\pm$ 2.0 $\pm$ 1.4	30	FRABETTI	93B E687	See FRABETTI 98
2459 $\pm$ 5 $\pm$ 30	56	<sup>2</sup> COTEUS	87 SPEC	$nA \approx 600$ GeV
2460 $\pm$ 25	82	BIAGI	83 SPEC	$\Sigma^-$ Be 135 GeV

<sup>1</sup> The systematic error was (wrongly) given the other way round in LESIAK 05; see the erratum.  
<sup>2</sup> Although COTEUS 87 claims to agree well with BIAGI 83 on the mass and width, there appears to be a discrepancy between the two experiments. BIAGI 83 sees a single peak (statd significance about 6 standard deviations) in the  $\Lambda K^- \pi^+ \pi^+$  mass spectrum. COTEUS 87 sees *two* peaks in the same spectrum, one at the  $\Xi_c^+$  mass, the other 75 MeV lower. The latter is attributed to  $\Xi_c^+ \rightarrow \Sigma^0 K^- \pi^+ \pi^+ \rightarrow (\Lambda \gamma) K^- \pi^+ \pi^+$ , with the  $\gamma$  unseen. The *combined* significance of the double peak is stated to be 5.5 standard deviations. But the absence of any trace of a lower peak in BIAGI 83 seems to us to throw into question the interpretation of the lower peak of COTEUS 87.

## $\Xi_c^+$ MEAN LIFE

VALUE ( $10^{-15}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>442<math>\pm</math> 26 OUR AVERAGE</b>				Error includes scale factor of 1.3. See the ideogram below.
503 $\pm$ 47 $\pm$ 18	250	MAHMOOD	02 CLE2	$e^+e^- \approx \Upsilon(4S)$
439 $\pm$ 22 $\pm$ 9	532	LINK	01D FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV
340 $\pm$ 70 50 $\pm$ 20	56	FRABETTI	98 E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
400 $\pm$ 180 120 $\pm$ 100	102	COTEUS	87 SPEC	$nA \approx 600$ GeV
480 $\pm$ 210 $\pm$ 200 150 $\pm$ 100	53	BIAGI	85c SPEC	$\Sigma^-$ Be 135 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
410 $\pm$ 110 80 $\pm$ 20	30	FRABETTI	93B E687	See FRABETTI 98
200 $\pm$ 110 60 $\pm$ 20	6	BARLAG	89c ACCM	$\pi^-$ ( $K^-$ ) Cu 230 GeV

WEIGHTED AVERAGE  
442 $\pm$ 26 (Error scaled by 1.3)



$\Xi_c^+$  mean life

## $\Xi_c^+$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
------	--------------------------------	------------------

No absolute branching fractions have been measured.  
The following are branching *ratios* relative to  $\Xi^- 2\pi^+$ .

### Cabibbo-favored ( $S = -2$ ) decays — relative to $\Xi^- 2\pi^+$

$\Gamma_1$	$p 2K_S^0$	0.087 $\pm$ 0.021	
$\Gamma_2$	$\Lambda \overline{K}^0 \pi^+$	—	
$\Gamma_3$	$\Sigma(1385)^+ \overline{K}^0$	[a] 1.0 $\pm$ 0.5	
$\Gamma_4$	$\Lambda K^- 2\pi^+$	0.323 $\pm$ 0.033	
$\Gamma_5$	$\Lambda \overline{K}^*(892)^0 \pi^+$	[a] <0.16	90%
$\Gamma_6$	$\Sigma(1385)^+ K^- \pi^+$	[a] <0.23	90%
$\Gamma_7$	$\Sigma^+ K^- \pi^+$	0.94 $\pm$ 0.10	
$\Gamma_8$	$\Sigma^+ \overline{K}^*(892)^0$	[a] 0.81 $\pm$ 0.15	
$\Gamma_9$	$\Sigma^0 K^- 2\pi^+$	0.27 $\pm$ 0.12	
$\Gamma_{10}$	$\Xi^0 \pi^+$	0.55 $\pm$ 0.16	
$\Gamma_{11}$	$\Xi^- 2\pi^+$	DEFINED AS 1	
$\Gamma_{12}$	$\Xi(1530)^0 \pi^+$	[a] <0.10	90%
$\Gamma_{13}$	$\Xi^0 \pi^+ \pi^0$	2.3 $\pm$ 0.7	
$\Gamma_{14}$	$\Xi^0 \pi^- 2\pi^+$	1.7 $\pm$ 0.5	
$\Gamma_{15}$	$\Xi^0 e^+ \nu_e$	2.3 $\pm$ 0.7 -0.8	
$\Gamma_{16}$	$\Omega^- K^+ \pi^+$	0.07 $\pm$ 0.04	

### Cabibbo-suppressed decays — relative to $\Xi^- 2\pi^+$

$\Gamma_{17}$	$p K^- \pi^+$	0.21 $\pm$ 0.04	
$\Gamma_{18}$	$p \overline{K}^*(892)^0$	[a] 0.116 $\pm$ 0.030	
$\Gamma_{19}$	$\Sigma^+ \pi^+ \pi^-$	0.48 $\pm$ 0.20	
$\Gamma_{20}$	$\Sigma^- 2\pi^+$	0.18 $\pm$ 0.09	
$\Gamma_{21}$	$\Sigma^+ K^+ K^-$	0.15 $\pm$ 0.06	
$\Gamma_{22}$	$\Sigma^+ \phi$	[a] <0.11	90%
$\Gamma_{23}$	$\Xi(1690)^0 K^+, \Xi^0 \rightarrow \Sigma^+ K^-$	<0.05	90%

[a] This branching fraction includes all the decay modes of the final-state resonance.

## $\Xi_c^+$ BRANCHING RATIOS

### Cabibbo-favored ( $S = -2$ ) decays

$\Gamma(p2K_S^0)/\Gamma(\Xi^-2\pi^+)$				$\Gamma_1/\Gamma_{11}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.087±0.016±0.014</b>	168 ± 27	LESLAK	05 BELL	$e^+e^-$ , $\Upsilon(4S)$

$\Gamma(\Sigma(1385)^+ \bar{K}^0)/\Gamma(\Xi^- 2\pi^+)$				$\Gamma_3/\Gamma_{11}$
Unseen decay modes of the $\Sigma(1385)^+$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.00<math>\pm</math>0.49<math>\pm</math>0.24</b>	20	LINK	03E FOCS	<1.72, 90% CL

$\Gamma(\Lambda K^- 2\pi^+)/\Gamma(\Xi^- 2\pi^+)$				$\Gamma_4/\Gamma_{11}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.323<math>\pm</math>0.033 OUR AVERAGE</b>				
0.32 $\pm$ 0.03 $\pm$ 0.02	1177 $\pm$ 55	LESIAK	05 BELL	$e^+e^-$ , $\Upsilon(4S)$
0.28 $\pm$ 0.06 $\pm$ 0.06	58	LINK	03E FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV
0.58 $\pm$ 0.16 $\pm$ 0.07	61	BERGFELD	96 CLE2	$e^+e^- \approx \Upsilon(4S)$

$\Gamma(\Lambda\overline{K}^*(892)^0\pi^+)/\Gamma(\Lambda K^-2\pi^+)$					$\Gamma_5/\Gamma_4$
Unseen decay modes of the $\overline{K}^*(892)^0$ are included.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.5</b>	90	BERGFELD	96 CLE2	$e^+e^- \approx \Upsilon(4S)$	

$\Gamma(\Sigma(1385)^+ K^- \pi^+)/\Gamma(\Lambda K^- 2\pi^+)$					$\Gamma_6/\Gamma_4$
Unseen decay modes of the $\Sigma(1385)^+$ are included.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<0.7	90	BERGFELD	96 CLE2	$e^+e^- \approx \Upsilon(4S)$	

$\Gamma(\Sigma^+ K^- \pi^+)/\Gamma(\Xi^- 2\pi^+)$				$\Gamma_7/\Gamma_{11}$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.94±0.10 OUR AVERAGE</b>				
0.91±0.11±0.04	251	LINK	03E FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV
0.92±0.20±0.07		<sup>3</sup> JUN	00 SELX	$\Sigma^-$ nucleus, 600 GeV
1.18±0.26±0.17	119	BERGFELD	96 CLE2	$e^+e^- \approx \Upsilon(4S)$

<sup>3</sup> This JUN 00 result is redundant with other results given below.

$\Gamma(\Sigma^+ \overline{K}^*(892)^0)/\Gamma(\Xi^- 2\pi^+)$				$\Gamma_8/\Gamma_{11}$
Unseen decay modes of the $\overline{K}^*(892)^0$ are included.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.81<math>\pm</math>0.15 OUR AVERAGE</b>				
0.78 $\pm$ 0.16 $\pm$ 0.06	119	LINK	03E FOCS	$\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV
0.92 $\pm$ 0.27 $\pm$ 0.14	61	BERGFELD	96 CLE2	$e^+e^- \approx \Upsilon(4S)$



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## Baryon Particle Listings

 $\Xi_c^+$ ,  $\Xi_c^0$  $\Gamma(\Sigma^0 K^- 2\pi^+)/\Gamma(\Lambda K^- 2\pi^+)$   $\Gamma_9/\Gamma_4$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.84±0.36</b>	47	<sup>4</sup> COTEUS	87	SPEC $nA \simeq 600$ GeV

<sup>4</sup> See, however, the note on the COTEUS 87  $\Xi_c^+$  mass measurement. $\Gamma(\Xi^0 \pi^+)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{10}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.55±0.13±0.09</b>	39	EDWARDS	96	CLE2 $e^+ e^- \approx \gamma(4S)$

 $\Gamma(\Xi^- 2\pi^+)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

seen	131	BERGFELD	96	CLE2 $e^+ e^- \approx \gamma(4S)$
seen	160	AVERY	95	CLE2 $e^+ e^- \approx \gamma(4S)$
seen	30	FRABETTI	93B	E687 $\gamma\text{Be}$ , $\overline{E}_\gamma = 220$ GeV
seen	30	ALBRECHT	90F	ARG $e^+ e^-$ at $\gamma(4S)$
seen	23	ALAM	89	CLEO $e^+ e^-$ 10.6 GeV

 $\Gamma(\Xi(1530)^0 \pi^+)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{12}/\Gamma_{11}$ Unseen decay modes of the  $\Xi(1530)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.1</b>	90	LINK	03E	FOCS $\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.2	90	BERGFELD	96	CLE2 $e^+ e^- \approx \gamma(4S)$
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 $\Gamma(\Xi^0 \pi^+ \pi^0)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{13}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.34±0.57±0.37</b>	81	EDWARDS	96	CLE2 $e^+ e^- \approx \gamma(4S)$

 $\Gamma(\Xi(1530)^0 \pi^+)/\Gamma(\Xi^0 \pi^+ \pi^0)$   $\Gamma_{12}/\Gamma_{13}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

<0.3	90	EDWARDS	96	CLE2 $e^+ e^- \approx \gamma(4S)$
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 $\Gamma(\Xi^0 \pi^- 2\pi^+)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{14}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.74±0.42±0.27</b>	57	EDWARDS	96	CLE2 $e^+ e^- \approx \gamma(4S)$

 $\Gamma(\Xi^0 e^+ \nu_e)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{15}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.3±0.6<sup>+0.3</sup><sub>-0.6</sub></b>	41	ALEXANDER	95B	CLE2 $e^+ e^- \approx \gamma(4S)$

 $\Gamma(\Omega^- K^+ \pi^+)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{16}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.07±0.03±0.03</b>	14	LINK	03E	FOCS < 0.12, 90% CL

## Cabibbo-suppressed decays

 $\Gamma(p K^- \pi^+)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{17}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21 ± 0.04 OUR AVERAGE</b>				

0.194±0.054	47 ± 11	VAZQUEZ-JA...08	SELX	$\Sigma^-$ nucleus, 600 GeV
0.234±0.047±0.022	202	LINK	01B	FOCS $\gamma$ nucleus

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.20 ± 0.04 ± 0.02	76	JUN	00	SELX See VAZQUEZ-JAUREGUI 08
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 $\Gamma(p \overline{K}^*(892)^0)/\Gamma(p K^- \pi^+)$   $\Gamma_{18}/\Gamma_{17}$ Unseen decay modes of the  $\overline{K}^*(892)^0$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.54±0.09±0.05</b>		LINK	01B	FOCS $\gamma$ nucleus

 $\Gamma(\Sigma^+ \pi^+ \pi^-)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{19}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.48±0.20</b>	21 ± 8	VAZQUEZ-JA...08	SELX	$\Sigma^-$ nucleus, 600 GeV

 $\Gamma(\Sigma^- 2\pi^+)/\Gamma(\Xi^- 2\pi^+)$   $\Gamma_{20}/\Gamma_{11}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.18±0.09</b>	10 ± 4	VAZQUEZ-JA...08	SELX	$\Sigma^-$ nucleus, 600 GeV

 $\Gamma(\Sigma^+ K^+ K^-)/\Gamma(\Sigma^+ K^- \pi^+)$   $\Gamma_{21}/\Gamma_7$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.16±0.06±0.01</b>	17	LINK	03E	FOCS $\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV

 $\Gamma(\Sigma^+ \phi)/\Gamma(\Sigma^+ K^- \pi^+)$   $\Gamma_{22}/\Gamma_7$ Unseen decay modes of the  $\phi$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.12</b>	90	LINK	03E	FOCS $\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV

 $\Gamma(\Xi(1690)^0 K^+ \times B(\Xi(1690)^0 \rightarrow \Sigma^+ K^-))/\Gamma(\Sigma^+ K^- \pi^+)$   $\Gamma_{23}/\Gamma_7$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.05</b>	90	LINK	03E	FOCS $\gamma$ nucleus, $\overline{E}_\gamma \approx 180$ GeV

 $\Xi_c^+$  REFERENCES

AALTONEN	14B	PR D89 072014	T. Aaltonen <i>et al.</i>	(CDF Collab.)
VAZQUEZ-JA...	08	PL B666 299	E. Vazquez-Jauregui <i>et al.</i>	(SELEX Collab.)
LESIAK	05	PL B605 237	T. Lesiak <i>et al.</i>	(BELLE Collab.)
Also		PL B617 198 (errat.)	T. Lesiak <i>et al.</i>	(BELLE Collab.)
LINK	03E	PL B571 139	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
MAHMOOD	02	PR D65 031102	A.H. Mahmood <i>et al.</i>	(CLEO Collab.)
LINK	01B	PL B512 277	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	01D	PL B523 53	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
JUN	00	PRL 84 1857	S.Y. Jun <i>et al.</i>	(FNAL SELEX Collab.)
FRABETTI	98	PL B427 211	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
BERGFELD	96	PL B365 431	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
EDWARDS	96	PL B373 261	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ALEXANDER	95B	PRL 74 3113	J. Alexander <i>et al.</i>	(CLEO Collab.)
Also		PRL 75 4155 (erratum)	J. Alexander <i>et al.</i>	(CLEO Collab.)
AVERY	95	PRL 75 4364	P. Avery <i>et al.</i>	(CLEO Collab.)
FRABETTI	93B	PRL 70 1381	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALBRECHT	90F	PL B247 121	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALAM	89	PL B226 401	M.S. Alam <i>et al.</i>	(CLEO Collab.)
BARLAG	89C	PL B233 522	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
COTEUS	87	PRL 59 1530	P. Coteus <i>et al.</i>	(FNAL E400 Collab.)
BIAGI	85C	PL 150B 230	S.F. Biagi <i>et al.</i>	(CERN WA62 Collab.)
BIAGI	83	PL 122B 455	S.F. Biagi <i>et al.</i>	(CERN WA62 Collab.)



$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+) \text{ Status: } ** *$$

According to the quark model, the  $\Xi_c^0$  (quark content  $dsc$ ) and  $\Xi_c^+$  form an isospin doublet, and the spin-parity ought to be  $J^P = 1/2^+$ . None of  $I$ ,  $J$ , or  $P$  has actually been measured.

 $\Xi_c^0$  MASSThe fit uses the  $\Xi_c^0$  and  $\Xi_c^+$  mass and mass-difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2470.85<sup>+0.28</sup><sub>-0.40</sub> OUR FIT</b>				

**2470.99<sup>+0.30</sup><sub>-0.50</sub> OUR AVERAGE**

2470.85 ± 0.24 ± 0.55	3.4k	AALTONEN	14B	CDF $p\overline{p}$ at 1.96 TeV
2471.0 ± 0.3 <sup>+0.2</sup> <sub>-1.4</sub>	8.6k	<sup>1</sup> LESIAK	05	BELL $e^+ e^-$ , $\gamma(4S)$
2470.0 ± 2.8 ± 2.6	85	FRABETTI	98B	E687 $\gamma\text{Be}$ , $\overline{E}_\gamma = 220$ GeV
2469 ± 2 ± 3	9	HENDERSON	92B	CLEO $\Omega^- K^+$
2472.1 ± 2.7 ± 1.6	54	ALBRECHT	90F	ARG $e^+ e^-$ at $\gamma(4S)$
2473.3 ± 1.9 ± 1.2	4	BARLAG	90	ACCM $\pi^- (K^-)$ Cu 230 GeV
2472 ± 3 ± 4	19	ALAM	89	CLEO $e^+ e^-$ 10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

2462.1 ± 3.1 ± 1.4	42	<sup>2</sup> FRABETTI	93C	E687 See FRABETTI 98B
2471 ± 3 ± 4	14	AVERY	89	CLEO See ALAM 89

<sup>1</sup> The systematic error was (wrongly) given the other way round in LESIAK 05.<sup>2</sup> The FRABETTI 93C mass is well below the other measurements. $\Xi_c^0 - \Xi_c^+$  MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.93±0.24 OUR FIT</b>				
<b>2.91±0.26 OUR AVERAGE</b>				
2.85 ± 0.30 ± 0.04	5.1/3.4k	AALTONEN	14B	CDF $p\overline{p}$ at 1.96 TeV
2.9 ± 0.5		LESIAK	05	BELL $e^+ e^-$ , $\gamma(4S)$
7.0 ± 4.5 ± 2.2		ALBRECHT	90F	ARG $e^+ e^-$ at $\gamma(4S)$
6.8 ± 3.3 ± 0.5		BARLAG	90	ACCM $\pi^- (K^-)$ Cu 230 GeV
5 ± 4 ± 1		ALAM	89	CLEO $\Xi_c^0 \rightarrow \Xi^- \pi^+$ , $\Xi_c^+ \rightarrow \Xi^- \pi^+ \pi^+$

 $\Xi_c^0$  MEAN LIFE

VALUE (10 <sup>-15</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>112<sup>+13</sup><sub>-10</sub> OUR AVERAGE</b>				
118 <sup>+14</sup> <sub>-12</sub> ± 5	110	LINK	02H	FOCS $\gamma$ nucleus, $\approx 180$ GeV
101 <sup>+25</sup> <sub>-17</sub> ± 5	42	FRABETTI	93C	E687 $\gamma\text{Be}$ , $\overline{E}_\gamma = 220$ GeV
82 <sup>+59</sup> <sub>-30</sub>	4	BARLAG	90	ACCM $\pi^- (K^-)$ Cu 230 GeV

 $\Xi_c^0$  DECAY MODES

No absolute branching fractions have been measured. Several measurements of ratios of fractions may be found in the Listings that follow.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
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# Baryon Particle Listings

$\Xi_c^0, \Xi_c^{'+}$

No absolute branching fractions have been measured. The following are branching <i>ratios</i> relative to $\Xi^- \pi^+$ .				
Cabibbo-favored (S = -2) decays — relative to $\Xi^- \pi^+$				
$\Gamma_1$	$\rho K^- K^- \pi^+$	0.34 ± 0.04		
$\Gamma_2$	$\rho K^- \bar{K}^*(892)^0$	0.21 ± 0.05		
$\Gamma_3$	$\rho K^- K^- \pi^+$ (no $\bar{K}^{*0}$ )	0.21 ± 0.04		
$\Gamma_4$	$\Lambda K_S^0$	0.210 ± 0.028		
$\Gamma_5$	$\Lambda K^- \pi^+$	1.07 ± 0.14		
$\Gamma_6$	$\Lambda \bar{K}^0 \pi^+ \pi^-$	seen		
$\Gamma_7$	$\Lambda K^- \pi^+ \pi^+ \pi^-$	seen		
$\Gamma_8$	$\Xi^- \pi^+$	DEFINED AS 1		
$\Gamma_9$	$\Xi^- \pi^+ \pi^+ \pi^-$	3.3 ± 1.4		
$\Gamma_{10}$	$\Omega^- K^+$	0.297 ± 0.024		
$\Gamma_{11}$	$\Xi^- e^+ \nu_e$	3.1 ± 1.1		
$\Gamma_{12}$	$\Xi^- \ell^+$ anything	1.0 ± 0.5		
Cabibbo-suppressed decays — relative to $\Xi^- \pi^+$				
$\Gamma_{13}$	$\Xi^- K^+$	0.028 ± 0.006		
$\Gamma_{14}$	$\Lambda K^+ K^-$ (no $\phi$ )	0.029 ± 0.007		
$\Gamma_{15}$	$\Lambda \phi$	0.034 ± 0.007		

## $\Xi_c^0$ BRANCHING RATIOS

Cabibbo-favored (S = -2) decays				
$\Gamma(pK^-K^-\pi^+)/\Gamma(\Xi^-\pi^+)$				$\Gamma_1/\Gamma_8$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.34±0.04 OUR AVERAGE</b>				
0.33±0.03±0.03	1908 ± 62	LESIAK	05	BELL $e^+e^-$ , $\Upsilon(4S)$
0.35±0.06±0.03	148 ± 18	DANKO	04	CLEO $e^+e^-$
$\Gamma(pK^-\bar{K}^*(892)^0)/\Gamma(\Xi^-\pi^+)$				$\Gamma_2/\Gamma_8$
Unseen decay modes of the $\bar{K}^*(892)^0$ are included.				
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.210±0.045±0.015</b>	DANKO	04	CLEO $e^+e^-$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	BARLAG	90	ACCM $\pi^- (K^-)$ Cu 230 GeV	
$\Gamma(pK^-K^-\pi^+ \text{ (no } \bar{K}^{*0}))/\Gamma(\Xi^-\pi^+)$				$\Gamma_3/\Gamma_8$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.21±0.04±0.02</b>	DANKO	04	CLEO $e^+e^-$	
$\Gamma(\Lambda K_S^0)/\Gamma(\Xi^-\pi^+)$				$\Gamma_4/\Gamma_8$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21±0.02±0.02</b>	465 ± 37	LESIAK	05	BELL $e^+e^-$ , $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
seen	7	ALBRECHT	95B	ARG $e^+e^- \approx 10.4$ GeV
$\Gamma(\Lambda K^-\pi^+)/\Gamma(\Xi^-\pi^+)$				$\Gamma_5/\Gamma_8$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.07±0.12±0.07</b>	2979 ± 211	LESIAK	05	BELL $e^+e^-$ , $\Upsilon(4S)$
$\Gamma(\Lambda \bar{K}^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	FRABETTI	98B	E687 $\gamma$ Be, $\bar{E}_\gamma = 220$ GeV	
$\Gamma(\Lambda K^-\pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$				$\Gamma_7/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	FRABETTI	98B	E687 $\gamma$ Be, $\bar{E}_\gamma = 220$ GeV	
$\Gamma(\Xi^-\pi^+)/\Gamma(\Xi^-\pi^+ \pi^+ \pi^-)$				$\Gamma_8/\Gamma_9$
VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.30±0.12±0.05</b>	ALBRECHT	90F	ARG $e^+e^-$ at $\Upsilon(4S)$	
$\Gamma(\Omega^-K^+)/\Gamma(\Xi^-\pi^+)$				$\Gamma_{10}/\Gamma_8$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.297±0.024 OUR AVERAGE</b>				
0.294±0.018±0.016	650	AUBERT,B	05M	BABR $e^+e^- \approx \Upsilon(4S)$
0.50±0.21±0.05	9	HENDERSON	92B	CLEO $e^+e^- \approx 10.6$ GeV
$\Gamma(\Xi^-e^+\nu_e)/\Gamma(\Xi^-\pi^+)$				$\Gamma_{11}/\Gamma_8$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.1±1.0<sup>+0.3</sup><sub>-0.5</sub></b>	54	ALEXANDER	95B	CLE2 $e^+e^- \approx \Upsilon(4S)$
$\Gamma(\Xi^-\ell^+ \text{ anything})/\Gamma(\Xi^-\pi^+)$				$\Gamma_{12}/\Gamma_8$
The ratio is for the <i>average</i> (not the sum) of the $\Xi^-e^+$ anything and $\Xi^-\mu^+$ anything modes.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.96±0.43±0.18</b>	18	ALBRECHT	93B	ARG $e^+e^- \approx 10.4$ GeV

$\Gamma(\Xi^- \ell^+ \text{ anything})/\Gamma(\Xi^- \pi^+ \pi^+ \pi^-)$			$\Gamma_{12}/\Gamma_9$	
The ratio is for the <i>average</i> (not the sum) of the $\Xi^- e^+$ anything and $\Xi^- \mu^+$ anything modes.				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.29±0.12±0.04</b>	18	ALBRECHT	93B	ARG $e^+e^- \approx 10.4$ GeV
<hr/> <b>Cabibbo-suppressed decays</b> <hr/>				
$\Gamma(\Xi^- K^+)/\Gamma(\Xi^- \pi^+)$			$\Gamma_{13}/\Gamma_8$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.75±0.51±0.25</b>	314 ± 58	CHISTOV	13	BELL $e^+e^- \approx \Upsilon(4S)$
$\Gamma(\Lambda K^+ K^- (\text{no } \phi))/\Gamma(\Xi^- \pi^+)$			$\Gamma_{14}/\Gamma_8$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.86±0.61±0.37</b>	510 ± 110	CHISTOV	13	BELL $e^+e^- \approx \Upsilon(4S)$
$\Gamma(\Lambda \phi)/\Gamma(\Xi^- \pi^+)$			$\Gamma_{15}/\Gamma_8$	
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.43±0.58±0.32</b>	316 ± 54	CHISTOV	13	BELL $e^+e^- \approx \Upsilon(4S)$

## $\Xi_c^0$ DECAY PARAMETERS

See the note on “Baryon Decay Parameters” in the neutron Listings.

$\alpha$ FOR $\Xi_c^0 \rightarrow \Xi^- \pi^+$				
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.56 ± 0.39 ± 0.10</b> <b>-0.09</b>	138	CHAN	01	CLE2 $e^+ e^- \approx \Upsilon(4S)$

## $\Xi_c^0$ REFERENCES

AALTONEN	14B	PR D89 072014	T. Aaltonen <i>et al.</i>	(CDF Collab.)
CHISTOV	13	PR D88 071103	R. Chistov <i>et al.</i>	(BELLE Collab.)
AUBERT,B	05M	PRL 95 142003	B. Aubert <i>et al.</i>	(BABAR Collab.)
LESIAK	05	PL B605 237	T. Lesiak <i>et al.</i>	(BELLE Collab.)
Also		PL B617 198 (errat.)	T. Lesiak <i>et al.</i>	(BELLE Collab.)
DANKO	04	PR D69 052004	I. Danko <i>et al.</i>	(CLEO Collab.)
LINK	02H	PL B541 211	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
CHAN	01	PR D63 111102	S. Chan <i>et al.</i>	(CLEO Collab.)
FRABETTI	98B	PL B426 403	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALBRECHT	95B	PL B342 397	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	95B	PRL 74 3113	J. Alexander <i>et al.</i>	(CLEO Collab.)
Also		PRL 75 4155 (erratum)	J. Alexander <i>et al.</i>	(CLEO Collab.)
ALBRECHT	93B	PL B303 368	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
FRABETTI	93C	PRL 70 2058	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
HENDERSON	92B	PL B283 161	S. Henderson <i>et al.</i>	(CLEO Collab.)
ALBRECHT	90F	PL B247 121	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARLAG	90	PL B236 495	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
ALAM	89	PL B226 401	M.S. Alam <i>et al.</i>	(CLEO Collab.)
AVERY	89	PRL 62 863	P. Avery <i>et al.</i>	(CLEO Collab.)



$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \*\*\*

The  $\Xi_c^{'+}$  and  $\Xi_c^0$  presumably complete the SU(3) sextet whose other members are the  $\Sigma_c^{++}$ ,  $\Sigma_c^+$ ,  $\Sigma_c^0$ , and  $\Omega_c^0$ ; see Fig. 3 in the note on Charmed Baryons just before the  $\Lambda_c^+$  Listings. The quantum numbers given above come from this presumption but have not been measured.

## $\Xi_c^{'+}$ MASS

The mass is obtained from the mass-difference measurement that follows.

VALUE (MeV)	DOCUMENT ID
<b>2575.7 ± 3.0 OUR FIT</b>	

## $\Xi_c^{'+} - \Xi_c^0$ MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>107.8 ± 3.0 OUR FIT</b>				
<b>107.8 ± 1.7 ± 2.5</b>	25	JESSOP	99	CLE2 $e^+ e^- \approx \Upsilon(4S)$

## $\Xi_c^{'+}$ DECAY MODES

The  $\Xi_c^{'+} - \Xi_c^0$  mass difference is too small for any strong decay to occur.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\Xi_c^+ \gamma$	seen

## $\Xi_c^{'+}$ REFERENCES

JESSOP	99	PRL 82 492	C.P. Jessop <i>et al.</i>	(CLEO Collab.)
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See key on page 601

Baryon Particle Listings

$\Xi_c^{\prime 0}$ ,  $\Xi_c(2645)$ ,  $\Xi_c(2790)$ ,  $\Xi_c(2815)$



$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \*\*\*

See the note in the Listing for the  $\Xi_c^{\prime +}$ , above.

$\Xi_c^0$  MASS

The mass is obtained from the mass-difference measurement that follows.

VALUE (MeV)	DOCUMENT ID
2577.9±2.9 OUR FIT	

$\Xi_c^0 - \Xi_c^0$  MASS DIFFERENCE

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
107.0±2.9 OUR FIT				
107.0±1.4±2.5	28	JESSOP	99 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

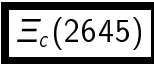
$\Xi_c^0$  DECAY MODES

The  $\Xi_c^0 - \Xi_c^0$  mass difference is too small for any strong decay to occur.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Xi_c^0 \gamma$	seen

$\Xi_c^0$  REFERENCES

JESSOP	99	PRL 82 492	C.P. Jessop <i>et al.</i>	(CLEO Collab.)
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$I(J^P) = \frac{1}{2}(\frac{3}{2}^+)$  Status: \*\*\*

The natural assignment is that this is the  $J^P = 3/2^+$  excitation of the  $\Xi_c$  in the same SU(4) multiplet as the  $\Delta(1232)$ , but the quantum numbers have not been measured.

$\Xi_c(2645)$  MASSES

The masses are obtained from the mass-difference measurements that follow.

$\Xi_c(2645)^+$ MASS	VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
2645.9±0.5 OUR FIT					Error includes scale factor of 1.1.
2645.6±0.2±0.6	578 ± 32	LESIAK	08 BELL	$e^+e^- \approx \mathcal{T}(4S)$	

$\Xi_c(2645)^0$ MASS	VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
2645.9±0.5 OUR FIT					
2645.7±0.2±0.6	611 ± 32	LESIAK	08 BELL	$e^+e^- \approx \mathcal{T}(4S)$	

$\Xi_c(2645) - \Xi_c$  MASS DIFFERENCES

$m_{\Xi_c(2645)^+} - m_{\Xi_c^0}$	VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
175.0±0.6 OUR FIT					Error includes scale factor of 1.1.
175.6±1.4 OUR AVERAGE					Error includes scale factor of 1.7.
177.1±0.5±1.1	47	FRABETTI	98B E687	$\gamma$ Be, $\overline{E}_\gamma = 220$ GeV	
174.3±0.5±1.0	34	GIBBONS	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$	

$m_{\Xi_c(2645)^0} - m_{\Xi_c^+}$	VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
178.0±0.6 OUR FIT					
178.2±0.5±1.0	55	AVERY	95 CLE2	$e^+e^- \approx \mathcal{T}(4S)$	

$\Xi_c(2645)^+ - \Xi_c(2645)^0$  MASS DIFFERENCE

$m_{\Xi_c(2645)^+} - m_{\Xi_c(2645)^0}$	VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
0.0±0.5 OUR FIT				
-0.1±0.3±0.6		LESIAK	08 BELL	$\approx 600$ evts each

$\Xi_c(2645)$  WIDTHS

$\Xi_c(2645)^+$ WIDTH	VALUE (MeV)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
2.6±0.2±0.4		3.7k		KATO	14 BELL	$e^+e^- \mathcal{T}(1S)$ to $\mathcal{T}(5S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<3.1	90			GIBBONS	96 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

$\Xi_c(2645)^0$  WIDTH

VALUE (MeV)	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
<5.5	90	55	AVERY	95 CLE2	$e^+e^- \approx \mathcal{T}(4S)$

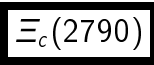
$\Xi_c(2645)$  DECAY MODES

$\Xi_c \pi$  is the only strong decay allowed to a  $\Xi_c$  resonance having this mass.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Xi_c^0 \pi^+$	seen
$\Gamma_2 \Xi_c^+ \pi^-$	seen

$\Xi_c(2645)$  REFERENCES

KATO	14	PR D89 052003	Y. Kato <i>et al.</i>	(BELLE Collab.)
LESIAK	08	PL B665 9	T. Lesiak <i>et al.</i>	(BELLE Collab.)
FRABETTI	98B	PL B426 403	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
GIBBONS	96	PRL 77 810	L.K. Gibbons <i>et al.</i>	(CLEO Collab.)
AVERY	95	PRL 75 4364	P. Avery <i>et al.</i>	(CLEO Collab.)



$I(J^P) = \frac{1}{2}(\frac{1}{2}^-)$  Status: \*\*\*

A peak seen in the  $\Xi_c' \pi$  mass spectrum. The simplest assignment, based on the mass, width, and decay mode, is that this belongs in the same SU(4) multiplet as the  $\Lambda(1405)$  and the  $\Lambda_c(2595)^+$ , but the spin and parity have not been measured.

$\Xi_c(2790)$  MASSES

The masses are obtained from the mass-difference measurements that follow.

$\Xi_c(2790)^+$ MASS	VALUE (MeV)	DOCUMENT ID
2789.1±3.2 OUR FIT		

$\Xi_c(2790)^0$ MASS	VALUE (MeV)	DOCUMENT ID
2791.9±3.3 OUR FIT		

$\Xi_c(2790) - \Xi_c$  MASS DIFFERENCES

$m_{\Xi_c(2790)^+} - m_{\Xi_c^0}$	VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
318.2±3.2 OUR FIT					
318.2±1.3±2.9	18	CSORNA	01 CLEO	$e^+e^- \approx \mathcal{T}(4S)$	

$m_{\Xi_c(2790)^0} - m_{\Xi_c^+}$	VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
324.0±3.3 OUR FIT					
324.0±1.3±3.0	14	CSORNA	01 CLEO	$e^+e^- \approx \mathcal{T}(4S)$	

$\Xi_c(2790)$  WIDTHS

$\Xi_c(2790)^+$ WIDTH	VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<15	90		CSORNA	01 CLEO	$e^+e^- \approx \mathcal{T}(4S)$

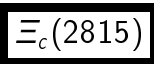
$\Xi_c(2790)^0$ WIDTH	VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<12	90		CSORNA	01 CLEO	$e^+e^- \approx \mathcal{T}(4S)$

$\Xi_c(2790)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Xi_c' \pi$	seen

$\Xi_c(2790)$  REFERENCES

CSORNA	01	PRL 86 4243	S.E. Csorna <i>et al.</i>	(CLEO Collab.)
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$I(J^P) = \frac{1}{2}(\frac{3}{2}^-)$  Status: \*\*\*

A narrow peak seen in the  $\Xi_c \pi \pi$  mass spectrum. The simplest assignment is that this belongs to the same SU(4) multiplet as the  $\Lambda(1520)$  and the  $\Lambda_c(2625)$ , but the spin and parity have not been measured.

Baryon Particle Listings

$\Xi_c(2815), \Xi_c(2930), \Xi_c(2970)$

$\Xi_c(2815)$  MASSES

The masses are obtained from the mass-difference measurements that follow.

$\Xi_c(2815)^+$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2816.6 \pm 0.9</math> OUR FIT</b>				
$2817.0 \pm 1.2^{+0.7}_{-0.8}$	73 $\pm$ 10	LESLIAK	08	BELL $e^+e^- \approx \tau(4S)$

$\Xi_c(2815)^0$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2819.6 \pm 1.2</math> OUR FIT</b>				
$2820.4 \pm 1.4^{+0.9}_{-1.0}$	48 $\pm$ 8	LESLIAK	08	BELL $e^+e^- \approx \tau(4S)$

$\Xi_c(2815) - \Xi_c$  MASS DIFFERENCES

$m_{\Xi_c(2815)^+} - m_{\Xi_c^+}$

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>348.7 \pm 0.9</math> OUR FIT</b>				
$348.6 \pm 0.6 \pm 1.0$	20	ALEXANDER	99B	CLE2 $e^+e^- \approx \tau(4S)$

$m_{\Xi_c(2815)^0} - m_{\Xi_c^0}$

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>348.8 \pm 1.2</math> OUR FIT</b>				
$347.2 \pm 0.7 \pm 2.0$	9	ALEXANDER	99B	CLE2 $e^+e^- \approx \tau(4S)$

$\Xi_c(2815)^+ - \Xi_c(2815)^0$  MASS DIFFERENCE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>-3.0 \pm 1.3</math> OUR FIT</b>			
$-3.4 \pm 1.9 \pm 0.9$	LESLIAK	08	BELL 73 & 48 events

$\Xi_c(2815)$  WIDTHS

$\Xi_c(2815)^+$  WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.5</b>	90	ALEXANDER	99B	CLE2 $e^+e^- \approx \tau(4S)$

$\Xi_c(2815)^0$  WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.5</b>	90	ALEXANDER	99B	CLE2 $e^+e^- \approx \tau(4S)$

$\Xi_c(2815)$  DECAY MODES

The  $\Xi_c \pi \pi$  modes are consistent with being entirely via  $\Xi_c(2645) \pi$ .

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Xi_c^+ \pi^+ \pi^-$	seen
$\Gamma_2 \Xi_c^0 \pi^+ \pi^-$	seen

$\Xi_c(2815)$  REFERENCES

LESLIAK	08	PL B665 9	T. Lesiak <i>et al.</i>	(BELLE Collab.)
ALEXANDER	99B	PRL 83 3390	J.P. Alexander <i>et al.</i>	(CLEO Collab.)

$\Xi_c(2930)$

$I(J^P) = ?(?^?)$  Status: \*

OMITTED FROM SUMMARY TABLE

A peak seen in the  $\Lambda_c^+ K^-$  mass projection of  $B^- \rightarrow \Lambda_c^+ \bar{\tau}_c^- K^-$  events.

$\Xi_c(2930)$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2931 \pm 3 \pm 5</math></b>	$\approx$ 34	AUBERT	08H	BABR $\tau(4S) \rightarrow B \bar{B}$

$\Xi_c(2930)$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>36 \pm 7 \pm 11</math></b>	$\approx$ 34	AUBERT	08H	BABR $\tau(4S) \rightarrow B \bar{B}$

$\Xi_c(2930)$  REFERENCES

AUBERT	08H	PR D77 031101	B. Aubert <i>et al.</i>	(BABAR Collab.)
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$\Xi_c(2970)$

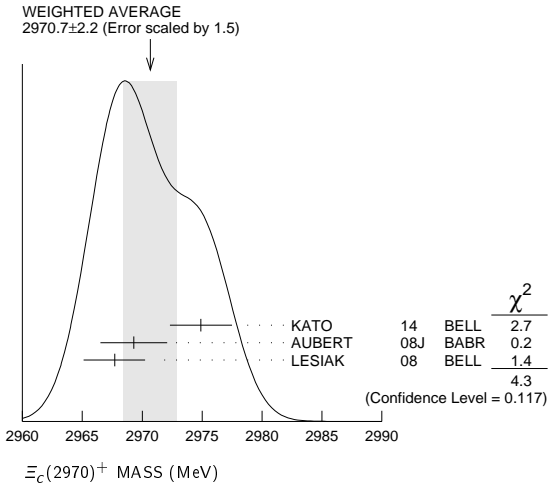
was  $\Xi_c(2980)$

$I(J^P) = \frac{1}{2}(?^?)$  Status: \*\*\*

$\Xi_c(2970)$  MASSES

$\Xi_c(2970)^+$  MASS

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2970.7 \pm 2.2</math> OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
$2974.9 \pm 1.5 \pm 2.1$	244 $\pm$ 39	KATO	14	BELL $e^+e^- \tau(1S)$ to $\tau(5S)$
$2969.3 \pm 2.2 \pm 1.7$	756 $\pm$ 206	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV
$2967.7 \pm 2.3^{+1.1}_{-1.2}$	78 $\pm$ 13	LESLIAK	08	BELL $e^+e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$2978.5 \pm 2.1 \pm 2.0$	405 $\pm$ 51	CHISTOV	06	BELL See KATO 14



$\Xi_c(2970)^0$  MASS

The evidence is statistically weaker for this charge state.

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>2968.0 \pm 2.6</math> OUR AVERAGE</b>				Error includes scale factor of 1.2.
$2972.9 \pm 4.4 \pm 1.6$	67 $\pm$ 44	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV
$2965.7 \pm 2.4^{+1.1}_{-1.2}$	57 $\pm$ 13	LESLIAK	08	BELL $e^+e^- \approx \tau(4S)$
$2977.1 \pm 8.8 \pm 3.5$	42 $\pm$ 24	CHISTOV	06	BELL $e^+e^- \approx \tau(4S)$

$\Xi_c(2970)$  WIDTHS

$\Xi_c(2970)^+$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>17.9 \pm 3.5</math> OUR AVERAGE</b>				
$14.8 \pm 2.5 \pm 4.1$	244 $\pm$ 39	KATO	14	BELL $e^+e^- \tau(1S)$ to $\tau(5S)$
27 $\pm$ 8 $\pm$ 3	756 $\pm$ 206	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV
18 $\pm$ 6 $\pm$ 3	78 $\pm$ 13	LESLIAK	08	BELL $e^+e^- \approx \tau(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$43.5 \pm 7.5 \pm 7.0$	405 $\pm$ 51	CHISTOV	06	BELL See KATO 14

$\Xi_c(2970)^0$  WIDTH

VALUE (MeV)	EVTs	DOCUMENT ID	TECN	COMMENT
<b><math>20 \pm 7</math> OUR AVERAGE</b>				Error includes scale factor of 1.3.
$31 \pm 7 \pm 8$	67 $\pm$ 44	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV
$15 \pm 6 \pm 3$	57 $\pm$ 13	LESLIAK	08	BELL $e^+e^- \approx \tau(4S)$

$\Xi_c(2970)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda_c^+ \bar{K} \pi$	seen
$\Gamma_2 \Sigma_c(2455) \bar{K}$	seen
$\Gamma_3 \Lambda_c^+ \bar{K}$	not seen
$\Gamma_4 \Xi_c 2\pi$	seen
$\Gamma_5 \Xi_c(2645) \pi$	seen

$\Xi_c(2970)$  BRANCHING RATIOS

$\Gamma(\Lambda_c^+ \bar{K} \pi)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	AUBERT	08J	BABR $e^+e^- \approx \Upsilon(4S)$
seen	CHISTOV	06	BELL $e^+e^- \approx \Upsilon(4S)$

See key on page 601

Baryon Particle Listings

$\Xi_c(2970)$ ,  $\Xi_c(3055)$ ,  $\Xi_c(3080)$ ,  $\Xi_c(3123)$ ,  $\Omega_c^0$

$\Gamma(\Sigma_c(2455)\bar{K})/\Gamma(\Lambda_c^+\bar{K}\pi)$				$\Gamma_2/\Gamma_1$	
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.55±0.07±0.13</b>	AUBERT	08J	BABR $e^+e^- \approx \mathcal{T}(4S)$		
$\Gamma(\Xi_c(2645)\pi)/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>seen</b>	LESIAK	08	BELL $e^+e^- \approx \mathcal{T}(4S)$		

$\Xi_c(2970)$  REFERENCES

KATO	14	PR D89 052003	Y. Kato <i>et al.</i>	(BELLE Collab.)
AUBERT	08J	PR D77 012002	B. Aubert <i>et al.</i>	(BABAR Collab.)
LESIAK	08	PL B665 9	T. Lesiak <i>et al.</i>	(BELLE Collab.)
CHISTOV	06	PRL 97 162001	R. Chistov <i>et al.</i>	(BELLE Collab.)

$\Xi_c(3055)$

$I(J^P) = ?(?^?)$  Status: \*\*\*

Seen in  $\Sigma_c(2455)^{++}K^- \rightarrow \Lambda_c^+K^-\pi^+$  with significances of 6.4 (BABAR) and 6.6 (BELLE) standard deviations.

$\Xi_c(3055)$  MASSES

$\Xi_c(3055)^+$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3055.1±1.7 OUR AVERAGE</b>		Error includes scale factor of 1.5.			
3058.1±1.0±2.1	199 ± 46	KATO	14	BELL $e^+e^- \mathcal{T}(1S)$ to $\mathcal{T}(5S)$	
3054.2±1.2±0.5	218 ± 95	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	

$\Xi_c(3055)$  WIDTHS

$\Xi_c(3055)^+$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>11 ±4 OUR AVERAGE</b>					
9.7±3.4± 3.3	199 ± 46	KATO	14	BELL $e^+e^- \mathcal{T}(1S)$ to $\mathcal{T}(5S)$	
17 ±6 ±11	218 ± 95	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	

$\Xi_c(3055)$  REFERENCES

KATO	14	PR D89 052003	Y. Kato <i>et al.</i>	(BELLE Collab.)
AUBERT	08J	PR D77 012002	B. Aubert <i>et al.</i>	(BABAR Collab.)

$\Xi_c(3080)$

$I(J^P) = \frac{1}{2}(?^?)$  Status: \*\*\*

A narrow peak seen in the  $\Lambda_c^+K^-\pi^+$  and  $\Lambda_c^+K_S^0\pi^-$  mass spectra.

$\Xi_c(3080)$  MASSES

$\Xi_c(3080)^+$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3076.94±0.28 OUR AVERAGE</b>					
3076.9 ±0.3 ±0.2	210 ± 30	KATO	14	BELL $e^+e^- \mathcal{T}(1S)$ to $\mathcal{T}(5S)$	
3077.0 ±0.4 ±0.2	403 ± 60	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
3076.7 ±0.9 ±0.5	326 ± 40	CHISTOV	06	BELL See KATO 14	

$\Xi_c(3080)^0$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3079.9±1.4 OUR AVERAGE</b>		Error includes scale factor of 1.3.			
3079.3±1.1±0.2	90 ± 27	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	
3082.8±1.8±1.5	67 ± 20	CHISTOV	06	BELL $e^+e^- \approx \mathcal{T}(4S)$	

$\Xi_c(3080)$  WIDTHS

$\Xi_c(3080)^+$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.3±1.5 OUR AVERAGE</b>		Error includes scale factor of 1.3.			
2.4±0.9±1.6	210 ± 30	KATO	14	BELL $e^+e^- \mathcal{T}(1S)$ to $\mathcal{T}(5S)$	
5.5±1.3±0.6	403 ± 60	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
6.2±1.2±0.8	326 ± 40	CHISTOV	06	BELL See KATO 14	

$\Xi_c(3080)^0$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>5.6±2.2 OUR AVERAGE</b>					
5.9±2.3±1.5	90 ± 27	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	
5.2±3.1±1.8	67 ± 20	CHISTOV	06	BELL $e^+e^- \approx \mathcal{T}(4S)$	

$\Xi_c(3080)$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda_c^+\bar{K}\pi$	seen
$\Gamma_2 \Sigma_c(2455)\bar{K}$	seen
$\Gamma_3 \Sigma_c(2455)\bar{K} + \Sigma_c(2520)\bar{K}$	seen
$\Gamma_4 \Lambda_c^+\bar{K}$	not seen
$\Gamma_5 \Lambda_c^+\bar{K}\pi^+\pi^-$	not seen

$\Xi_c(3080)$  BRANCHING RATIOS

$\Gamma(\Sigma_c(2455)\bar{K})/\Gamma(\Lambda_c^+\bar{K}\pi)$				$\Gamma_2/\Gamma_1$	
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.45±0.06 OUR AVERAGE</b>					
0.45±0.05±0.05	AUBERT	08J	BABR in $\Lambda_c^+K^-\pi^+$		
0.44±0.12±0.07	AUBERT	08J	BABR in $\Lambda_c^+K_S^0\pi^-$		
$[\Gamma(\Sigma_c(2455)\bar{K}) + \Gamma(\Sigma_c(2520)\bar{K})]/\Gamma(\Lambda_c^+\bar{K}\pi)$				$\Gamma_3/\Gamma_1$	
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.89±0.12 OUR AVERAGE</b>					
0.95±0.14±0.06	AUBERT	08J	BABR in $\Lambda_c^+K^-\pi^+$		
0.78±0.21±0.05	AUBERT	08J	BABR in $\Lambda_c^+K_S^0\pi^-$		

$\Xi_c(3080)$  REFERENCES

KATO	14	PR D89 052003	Y. Kato <i>et al.</i>	(BELLE Collab.)
AUBERT	08J	PR D77 012002	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHISTOV	06	PRL 97 162001	R. Chistov <i>et al.</i>	(BELLE Collab.)

$\Xi_c(3123)$

$I(J^P) = ?(?^?)$  Status: \*

OMITTED FROM SUMMARY TABLE

A peak in the  $\Sigma_c(2520)^{++}K^- \rightarrow \Lambda_c^+K^-\pi^+$  mass spectrum with a significance of 3.6 standard deviations. KATO 14 finds no evidence for this state.

$\Xi_c(3123)$  MASSES

$\Xi_c(3123)^+$ MASS					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>3122.9±1.3±0.3</b>	101 ± 35	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	

$\Xi_c(3123)$  WIDTHS

$\Xi_c(3123)^+$ WIDTH					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.4±3.4±1.7</b>	101 ± 35	AUBERT	08J	BABR $e^+e^- \approx 10.58$ GeV	

$\Xi_c(3123)$  REFERENCES

KATO	14	PR D89 052003	Y. Kato <i>et al.</i>	(BELLE Collab.)
AUBERT	08J	PR D77 012002	B. Aubert <i>et al.</i>	(BABAR Collab.)

$\Omega_c^0$

$I(J^P) = 0(\frac{1}{2}^+)$  Status: \*\*\*

The quantum numbers have not been measured, but are simply assigned in accord with the quark model, in which the  $\Omega_c^0$  is the ssc ground state.

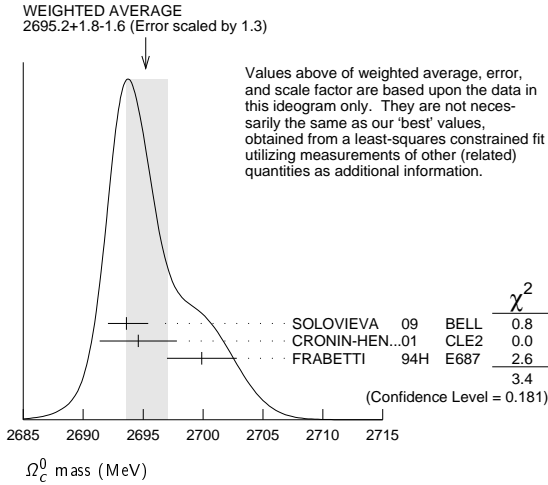
$\Omega_c^0$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2695.2± 1.7 OUR FIT</b>		Error includes scale factor of 1.3.		
<b>2695.2± 1.8 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
2693.6± 0.3±1.8	725 ± 45	SOLOVIEVA	09	BELL $\Omega^-\pi^+$ in $e^+e^- \rightarrow \mathcal{T}(4S)$
2694.6± 2.6±1.9	40	<sup>1</sup> CRONIN-HEN.	.01	CLE2 $e^+e^- \approx 10.6$ GeV
2699.9± 1.5±2.5	42	<sup>2</sup> FRABETTI	94H	E687 $\gamma$ Be, $\overline{E}_\gamma = 221$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2705.9± 3.3±2.0	10	<sup>3</sup> FRABETTI	93	E687 $\gamma$ Be, $\overline{E}_\gamma = 221$ GeV
2719.0± 7.0±2.5	11	<sup>4</sup> ALBRECHT	92H	ARG $e^+e^- \approx 10.6$ GeV
2740 ±20	3	BIAGI	85B	SPEC $\Sigma^-$ Be 135 GeV/c

# Baryon Particle Listings

## $\Omega_c^0, \Omega_c(2770)^0$

- <sup>1</sup> CRONIN-HENNESSY 01 sees  $40.4 \pm 9.0$  events in a sum over five channels.  
<sup>2</sup> FRABETTI 94H claims a signal of  $42.5 \pm 8.8 \Sigma^+ K^- K^- \pi^+$  events. The background is about 24 events.  
<sup>3</sup> FRABETTI 93 claims a signal of  $10.3 \pm 3.9 \Omega^- \pi^+$  events above a background of 5.8 events.  
<sup>4</sup> ALBRECHT 92H claims a signal of  $11.5 \pm 4.3 \Xi^- K^- \pi^+ \pi^+$  events. The background is about 5 events.



### $\Omega_c^0$ MEAN LIFE

VALUE (10 <sup>-15</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>69±12 OUR AVERAGE</b>				
72±11±11	64	LINK	03c	FOCS $\Omega^- \pi^+, \Xi^- K^- \pi^+ \pi^+$
55 <sup>+13+18</sup> <sub>-11-23</sub>	86	ADAMOVICH	95B	WA89 $\Omega^- \pi^- \pi^+ \pi^+, \Xi^- K^- \pi^+ \pi^+$
86 <sup>+27</sup> <sub>-20</sub> ±28	25	FRABETTI	95D	E687 $\Sigma^+ K^- K^- \pi^+$

### $\Omega_c^0$ DECAY MODES

No absolute branching fractions have been measured.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Sigma^+ K^- K^- \pi^+$	seen
$\Gamma_2 \quad \Xi^0 K^- \pi^+$	seen
$\Gamma_3 \quad \Xi^- K^- \pi^+ \pi^+$	seen
$\Gamma_4 \quad \Omega^- e^+ \nu_e$	seen
$\Gamma_5 \quad \Omega^- \pi^+$	seen
$\Gamma_6 \quad \Omega^- \pi^+ \pi^0$	seen
$\Gamma_7 \quad \Omega^- \pi^- \pi^+ \pi^+$	seen

### $\Omega_c^0$ BRANCHING RATIOS

$\Gamma(\Sigma^+ K^- K^- \pi^+)/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
seen	42	FRABETTI	94H	E687	$\gamma\text{Be}, \overline{E}_\gamma = 221 \text{ GeV}$	

$\Gamma(\Sigma^+ K^- K^- \pi^+)/\Gamma(\Omega^- \pi^+)$					$\Gamma_1/\Gamma_5$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<4.8	90	CRONIN-HEN..01	CLE2	$e^+ e^- \approx 10.6 \text{ GeV}$	

$\Gamma(\Xi^0 K^- \pi^+)/\Gamma(\Omega^- \pi^+)$				$\Gamma_2/\Gamma_5$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.0±2.5±0.4</b>	9	CRONIN-HEN..01	CLE2	$e^+ e^- \approx 10.6 \text{ GeV}$

$\Gamma(\Xi^- K^- \pi^+ \pi^+)/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	11	ALBRECHT	92H	ARG $e^+ e^- \approx 10.6 \text{ GeV}$
seen	3	BIAGI	85B	SPEC $\Sigma^- \text{Be } 135 \text{ GeV}/c$

$(\Xi^- K^- \pi^+ \pi^+)/\Gamma(\Omega^- \pi^+)$						$\Gamma_3/\Gamma_5$	
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>0.46±0.13±0.03</b>		45 ± 12	AUBERT	07AH	BABR	$e^+ e^- \approx \mathcal{T}(45)$	
• • • We do not use the following data for averages, fits, limits, etc. • • •							
1.6 ±1.1 ±0.4		7	CRONIN-HEN..01	CLE2		$e^+ e^- \approx 10.6 \text{ GeV}$	
<2.8	90		FRABETTI	93	E687	$\gamma \text{Be}, \overline{E}_\gamma = 221 \text{ GeV}$	

$\Gamma(\Omega^- \pi^+)/\Gamma(\Omega^- e^+ \nu_e)$					$\Gamma_5/\Gamma_4$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.41±0.19±0.04</b>	11	AMMAR	02	CLE2	$e^+ e^- \approx \mathcal{T}(4S)$

$\Gamma(\Omega^- \pi^+ \pi^0)/\Gamma(\Omega^- \pi^+)$				$\Gamma_6/\Gamma_5$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.27±0.31±0.11</b>	64 ± 15	AUBERT	07AH	BABR $e^+e^- \approx \tau(45)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.2 ±2.2 ±0.9	12	CRONIN-HEN..01	CLE2	$e^+e^- \approx 10.6 \text{ GeV}$

$\Gamma(\Omega^-\pi^-\pi^+\pi^+)/\Gamma(\Omega^-\pi^+)$					$\Gamma_7/\Gamma_5$
VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.28±0.09±0.01</b>		25 ± 8	AUBERT	07AH	BABR $e^+e^- \approx \mathcal{T}(45)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.56	90		CRONIN-HEN..01	CLE2	$e^+e^- \approx 10.6$ GeV
seen			ADAMOVICH	95B	WA89 $\Sigma^-$ 340 GeV
<1.6	90		FRABETTI	93	E687 $\gamma$ Be, $\overline{E}_\gamma = 221$ GeV

### $\Omega_c^0$ REFERENCES

SOLOVIEVA	09	PL B672 1	E. Solovieva <i>et al.</i>	(BELLE Collab.)
AUBERT	07AH	PRL 99 062001	B. Aubert <i>et al.</i>	(BABAR Collab.)
LINK	03c	PL B561 41	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
AMMAR	02	PRL 89 171803	R. Ammar <i>et al.</i>	(CLEO Collab.)
CRONIN-HEN..01	PRL	86 3730	D. Cronin-Hennessy <i>et al.</i>	(CLEO Collab.)
ADAMOVICH	95B	PL B358 151	M.I. Adamovich <i>et al.</i>	(CERN WA89 Collab.)
FRABETTI	95D	PL B357 678	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	94H	PL B338 106	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	93	PL B300 190	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALBRECHT	92H	PL B288 367	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BIAGI	85B	ZPHY C28 175	S.F. Biagi <i>et al.</i>	(CERN WA62 Collab.)

## $\Omega_c(2770)^0$

$I(J^P) = 0(\frac{3}{2}^+)$  Status: \* \* \*

The natural assignment is that this goes with the  $\Sigma_c(2520)$  and  $\Xi_c(2645)$  to complete the lowest mass  $J^P = \frac{3}{2}^+$  SU(3) sextet, part of the SU(4) 20-plet that includes the  $\Delta(1232)$ . But  $J$  and  $P$  have not been measured.

### $\Omega_c(2770)^0$ MASS

The mass is obtained from the mass-difference measurement that follows.

VALUE (MeV)	DOCUMENT ID
<b>2765.9±2.0 OUR FIT</b>	Error includes scale factor of 1.2.

### $\Omega_c(2770)^0 - \Omega_c^0$ MASS DIFFERENCE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>70.7<sup>+0.8</sup><sub>-0.9</sub> OUR FIT</b>				
<b>70.7<sup>+0.8</sup><sub>-1.0</sub> OUR AVERAGE</b>				
70.7±0.9 <sup>+0.1</sup> <sub>-0.9</sub>	54±9	SOLOVIEVA	09	BELL $\Omega_c^0 \gamma$ in $e^+ e^- \rightarrow \mathcal{T}(45)$
70.8±1.0±1.1	105±22	AUBERT,BE	06i	BABR $e^+ e^- \approx \mathcal{T}(45)$

### $\Omega_c(2770)^0$ DECAY MODES

The  $\Omega_c(2770)^0 - \Omega_c^0$  mass difference is too small for any strong decay to occur.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Omega_c^0 \gamma$	presumably 100%

### $\Omega_c(2770)^0$ REFERENCES

SOLOVIEVA	09	PL B672 1	E. Solovieva <i>et al.</i>	(BELLE Collab.)
AUBERT,BE	06i	PRL 97 232001	B. Aubert <i>et al.</i>	(BABAR Collab.)

See key on page 601

Baryon Particle Listings

$\Xi_{cc}^{++}$

DOUBLY CHARMED BARYONS  
(C = +2)  
 $\Xi_{cc}^{++} = u c c, \Xi_{cc}^{+} = d c c, \Omega_{cc}^{+} = s c c$

$\Xi_{cc}^{++}$

$I(J^P) = ?(?^?)$  Status: \*

OMITTED FROM SUMMARY TABLE

This would presumably be an isospin-1/2 particle, a  $ccu \Xi_{cc}^{++}$  and a  $ccd \Xi_{cc}^{+}$ . However, opposed to the evidence cited below, the BABAR experiment has found no evidence for a  $\Xi_{cc}^{+}$  in a search in  $\Lambda_c^{+} K^{-} \pi^{+}$  and  $\Xi_c^0 \pi^{+}$  modes, and no evidence of a  $\Xi_{cc}^{++}$  in  $\Lambda_c^{+} K^{-} \pi^{+} \pi^{+}$  and  $\Xi_c^0 \pi^{+} \pi^{+}$  modes (AUBERT,B 06D). Nor have the BELLE (CHISTOV 06, KATO 14) or LHCb (AAIJ 13CD) experiments found any evidence for this state.

$\Xi_{cc}^{+}$  MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3518.9±0.9 OUR AVERAGE</b>				
3518 ±3	6	<sup>1</sup> OCHERASHVI..05	SELX	$\Sigma^{-}$ nucleus $\approx$ 600 GeV
3519 ±1	16	<sup>2</sup> MATTSON 02	SELX	$\Sigma^{-}$ nucleus $\approx$ 600 GeV

<sup>1</sup> OCHERASHVILI 05 claims “an excess of 5.62 events over ...  $1.38 \pm 0.13$  events” for a significance of  $4.8 \sigma$  in  $p D^{+} K^{-}$  events.  
<sup>2</sup> MATTSON 02 claims “an excess of 15.9 events over an expected background of  $6.1 \pm 0.5$  events, a statistical significance of  $6.3 \sigma$ ” in the  $\Lambda_c^{+} K^{-} \pi^{+}$  invariant-mass spectrum. The probability that the peak is a fluctuation increases from  $1.0 \times 10^{-6}$  to  $1.1 \times 10^{-4}$  when the number of bins searched is considered.

$\Xi_{cc}^{+}$  MEAN LIFE

VALUE ( $10^{-15}$ s)	CL%	DOCUMENT ID	TECN	COMMENT
<33	90	MATTSON	02	SELX $\Sigma^{-}$ nucleus, $\approx$ 600 GeV

$\Xi_{cc}^{+}$  DECAY MODES

Mode	
$\Gamma_1$	$\Lambda_c^{+} K^{-} \pi^{+}$
$\Gamma_2$	$p D^{+} K^{-}$

$\Gamma(p D^{+} K^{-})/\Gamma(\Lambda_c^{+} K^{-} \pi^{+})$				$\Gamma_2/\Gamma_1$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.36±0.21</b>	6	UCHERASHVI..05	SELX	$\Sigma^{-}$ $\approx$ 600 GeV

$\Xi_{cc}^{+}$  REFERENCES

KATO	14	PR D89 052003	Y. Kato <i>et al.</i>	(BELLE Collab.)
AAIJ	13CD	JHEP 1312 090	R. Aaij <i>et al.</i>	(LHCb Collab.)
AUBERT,B	06D	PR D74 011103	B. Aubert <i>et al.</i>	(BABAR Collab.)
CHISTOV	06	PRL 97 162001	R. Chistov <i>et al.</i>	(BELLE Collab.)
UCHERASHVI..05	PL	B628 18	A. Ocherashvili <i>et al.</i>	(FNAL SELEX Collab.)
MATTSON	02	PRL 89 112001	M. Mattson <i>et al.</i>	(FNAL SELEX Collab.)

# Baryon Particle Listings

$\Lambda_b^0$

## BOTTOM BARYONS ( $B = -1$ )

$\Lambda_b^0 = udb, \Xi_b^0 = usb, \Xi_b^- = dsb, \Omega_b^- = ssb$

$\Lambda_b^0$

$I(J^P) = 0(\frac{1}{2}^+)$  Status: \*\* \*

In the quark model, a  $\Lambda_b^0$  is an isospin-0  $u d b$  state. The lowest  $\Lambda_b^0$  ought to have  $J^P = 1/2^+$ . None of  $I, J$ , or  $P$  have actually been measured.

### $\Lambda_b^0$ MASS

$m_{\Lambda_b^0}$		EVTS	DOCUMENT ID	TECN	COMMENT
VALUE (MeV)					
<b>5619.51 ± 0.23 OUR AVERAGE</b>					
5619.30 ± 0.34	1	AAIJ	14AA	LHCB	$p\bar{p}$ at 7 TeV
5620.15 ± 0.31 ± 0.47	2	AALTONEN	14B	CDF	$p\bar{p}$ at 1.96 TeV
5619.7 ± 0.7 ± 1.1	2	AAD	13U	ATLS	$p\bar{p}$ at 7 TeV
5619.44 ± 0.13 ± 0.38	2	AAIJ	13AV	LHCB	$p\bar{p}$ at 7 TeV
5621 ± 4 ± 3	3	ABE	97B	CDF	$p\bar{p}$ at 1.8 TeV
5668 ± 16 ± 8	4	ABREU	96N	DLPH	$e^+e^- \rightarrow Z$
5614 ± 21 ± 4	4	BUSKULIC	96L	ALEP	$e^+e^- \rightarrow Z$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
5619.19 ± 0.70 ± 0.30	2	AAIJ	12E	LHCB	Repl. by AAIJ 13AV
5619.7 ± 1.2 ± 1.2	5	ACOSTA	06	CDF	Repl. by AALTO-NEN 14B
not seen	6	ABE	93B	CDF	Repl. by ABE 97B
5640 ± 50 ± 30	7	ALBAJAR	91E	UA1	$p\bar{p}$ 630 GeV
5640 +100 -210		BARI	91	SFM	$\Lambda_b^0 \rightarrow \rho D^0 \pi^-$
5650 +150 -200	90	BARI	91	SFM	$\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^+ \pi^- \pi^-$

- 1 Uses exclusively reconstructed final states  $\Lambda_b^0 \rightarrow \Lambda_c^+ D_s^-, \Lambda_c^+ D^-$  and  $\bar{B}^0 \rightarrow D^+ D_s^-$  decays. The uncertainty includes both statistical and systematic contributions.
- 2 Uses  $\Lambda_b^0 \rightarrow J/\psi \Lambda$  fully reconstructed decays.
- 3 ABE 97B observed 38 events with a background of  $18 \pm 1.6$  events in the mass range 5.60–5.65 GeV/ $c^2$ , a significance of  $> 3.4$  standard deviations.
- 4 Uses 4 fully reconstructed  $\Lambda_b$  events.
- 5 Uses exclusively reconstructed final states containing a  $J/\psi \rightarrow \mu^+ \mu^-$  decays.
- 6 ABE 93B states that, based on the signal claimed by ALBAJAR 91E, CDF should have found  $30 \pm 23 \Lambda_b^0 \rightarrow J/\psi(1S) \Lambda$  events. Instead, CDF found not more than 2 events.
- 7 ALBAJAR 91E claims  $16 \pm 5$  events above a background of  $9 \pm 1$  events, a significance of about 5 standard deviations.

### $m_{\Lambda_b^0} - m_{B^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>339.2 ± 1.4 ± 0.1</b>	1	ACOSTA	06 CDF $p\bar{p}$ at 1.96 TeV

- 1 Uses exclusively reconstructed final states containing  $J/\psi \rightarrow \mu^+ \mu^-$  decays.

### $m_{\Lambda_b^0} - m_{B^+}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>339.72 ± 0.28 OUR AVERAGE</b>			
339.72 ± 0.24 ± 0.18	1	AAIJ	14AA LHCB $p\bar{p}$ at 7 TeV
339.71 ± 0.71 ± 0.09	2	AAIJ	12E LHCB $p\bar{p}$ at 7 TeV

- 1 Uses exclusively reconstructed final states  $\Lambda_b^0 \rightarrow \Lambda_c^+ D_s^-, \Lambda_c^+ D^-$  and  $\bar{B}^0 \rightarrow D^+ D_s^-$  decays.
- 2 Uses exclusively reconstructed final states containing  $J/\psi \rightarrow \mu^+ \mu^-$  decays.

### $\Lambda_b^0$ MEAN LIFE

See  $b$ -baryon Admixture section for data on  $b$ -baryon mean life average over species of  $b$ -baryon particles.

“OUR EVALUATION” is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

VALUE ( $10^{-12}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.466 ± 0.010 OUR EVALUATION</b>				
1.415 ± 0.027 ± 0.006	1	AAIJ	14E	LHCB $p\bar{p}$ at 7 TeV
1.479 ± 0.009 ± 0.010	2	AAIJ	14U	LHCB $p\bar{p}$ at 7, 8 TeV
1.565 ± 0.035 ± 0.020	1	AALTONEN	14B	CDF $p\bar{p}$ at 1.96 TeV
1.449 ± 0.036 ± 0.017	1	AAD	13U	ATLS $p\bar{p}$ at 7 TeV
1.503 ± 0.052 ± 0.031	1	CHATRCHYAN	13AC	CMS $p\bar{p}$ at 7 TeV
1.303 ± 0.075 ± 0.035	1	ABAZOV	12U	D0 $p\bar{p}$ at 1.96 TeV
1.401 ± 0.046 ± 0.035	3	AALTONEN	10B	CDF $p\bar{p}$ at 1.96 TeV
1.290 +0.119 -0.087 -0.110 -0.091	4	ABAZOV	07U	D0 $p\bar{p}$ at 1.96 TeV
1.11 +0.19 -0.18 ± 0.05	5	ABREU	99W	DLPH $e^+e^- \rightarrow Z$
1.29 +0.24 -0.22 ± 0.06	5	ACKERSTAFF	98G	OPAL $e^+e^- \rightarrow Z$
1.21 ± 0.11	5	BARATE	98D	ALEP $e^+e^- \rightarrow Z$
1.32 ± 0.15 ± 0.07	6	ABE	96M	CDF $p\bar{p}$ at 1.8 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.482 ± 0.018 ± 0.012	7	AAIJ	13BB	LHCB	Repl. by AAIJ 14U
1.537 ± 0.045 ± 0.014	1	AALTONEN	11	CDF	Repl. by AALTONEN 14B
1.218 +0.130 -0.115 ± 0.042	1	ABAZOV	07S	D0	Repl. by ABAZOV 12U
1.593 +0.083 -0.078 ± 0.033	1	ABULENCIA	07A	CDF	Repl. by AALTONEN 11
1.22 +0.22 -0.18 ± 0.04	1	ABAZOV	05c	D0	Repl. by ABAZOV 07S
1.19 +0.21 -0.18 +0.07 -0.08		ABREU	96D	DLPH	Repl. by ABREU 99W
1.14 +0.22 -0.19 ± 0.07	69	AKERS	95K	OPAL	Repl. by ACKERSTAFF 98G
1.02 +0.23 -0.18 ± 0.06	44	BUSKULIC	95L	ALEP	Repl. by BARATE 98D

- 1 Measured mean life using fully reconstructed  $\Lambda_b^0 \rightarrow J/\psi \Lambda$  decays.
- 2 Used  $\Lambda_b^0 \rightarrow J/\psi p K^-$  decays.
- 3 Measured mean life using fully reconstructed  $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$  decays.
- 4 Measured using semileptonic decays  $\Lambda_b^0 \rightarrow \Lambda_c^+ \mu \nu X$  and  $\Lambda_c^+ \rightarrow K_S^0 p$ .
- 5 Measured using  $\Lambda_c \ell^-$  and  $\Lambda \ell^+ \ell^-$ .
- 6 Excess  $\Lambda_c \ell^-$ , decay lengths.
- 7 Measured the lifetime ratio of decays  $\Lambda_b^0 \rightarrow J/\psi p K^-$  to  $B^0 \rightarrow J/\psi \pi^+ K^-$  to be  $0.976 \pm 0.012 \pm 0.006$  with  $\tau_{B^0} = 1.519 \pm 0.007$  ps.

### $\tau_{\Lambda_b^0}/\tau_{B^0}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.940 ± 0.035 ± 0.006</b>	1	AAIJ	14E LHCB $p\bar{p}$ at 7 TeV

- 1 Measured using  $\Lambda_b^0 \rightarrow J/\psi \Lambda$  decays.

### $\tau_{\Lambda_b^0}/\tau_{B^0}$ MEAN LIFE RATIO

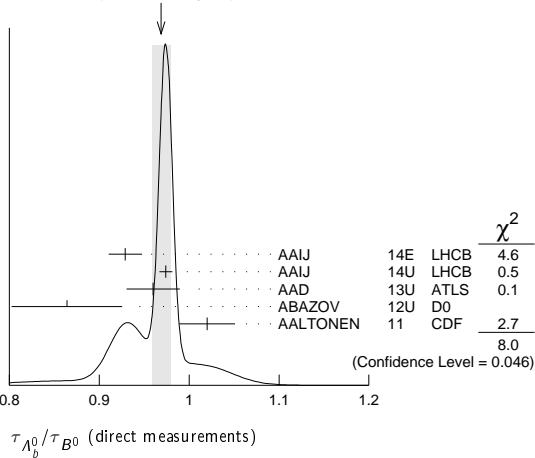
#### $\tau_{\Lambda_b^0}/\tau_{B^0}$ (direct measurements)

“OUR EVALUATION” has been obtained by the Heavy Flavor Averaging Group (HFAG) by including both  $B^0$  and  $B^+$  decays.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.964 ± 0.007 OUR EVALUATION</b>			
0.969 ± 0.010 OUR AVERAGE			Error includes scale factor of 1.6. See the ideogram below.
0.929 ± 0.018 ± 0.004	1	AAIJ	14E LHCB $p\bar{p}$ at 7 TeV
0.974 ± 0.006 ± 0.004	2	AAIJ	14U LHCB $p\bar{p}$ at 7, 8 TeV
0.960 ± 0.025 ± 0.016	3	AAD	13U ATLS $p\bar{p}$ at 7 TeV
0.864 ± 0.052 ± 0.033	4,5	ABAZOV	12U D0 $p\bar{p}$ at 1.96 TeV
1.020 ± 0.030 ± 0.008	4	AALTONEN	11 CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.976 ± 0.012 ± 0.006	6	AAIJ	13BB LHCB
0.811 +0.096 -0.087 ± 0.034	4,5	ABAZOV	07S D0
1.041 ± 0.057	7	ABULENCIA	07A CDF
0.87 +0.17 -0.14 ± 0.03	7	ABAZOV	05c D0

- 1 Measured using  $\Lambda_b^0 \rightarrow J/\psi \Lambda$  and  $B^0 \rightarrow J/\psi K^{*0}$  decays.
- 2 Used  $\Lambda_b^0 \rightarrow J/\psi p K^-$  and  $B^0 \rightarrow J/\psi K^{*0}(892)^0$  decays.
- 3 Measured with  $\Lambda_b^0 \rightarrow J/\psi(\mu^+ \mu^-) \Lambda^0(p\pi^-)$  decays.
- 4 Uses fully reconstructed  $\Lambda_b \rightarrow J/\psi \Lambda$  decays.
- 5 Uses  $B^0 \rightarrow J/\psi K_S^0$  decays for denominator.
- 6 Measures  $1/\tau_{\Lambda_b^0} - 1/\tau_{B^0}$  and uses  $\tau_{B^0} = 1.519 \pm 0.007$  ps to extract lifetime ratio.
- 7 Measured mean life ratio using fully reconstructed decays.

WEIGHTED AVERAGE  
0.969±0.010 (Error scaled by 1.6)





$\Lambda_b^0$  DECAY MODES

The branching fractions  $B(b\text{-baryon} \rightarrow \Lambda \ell^- \bar{\nu}_\ell \text{anything})$  and  $B(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})$  are not pure measurements because the underlying measured products of these with  $B(b \rightarrow b\text{-baryon})$  were used to determine  $B(b \rightarrow b\text{-baryon})$ , as described in the note "Production and Decay of  $b$ -Flavored Hadrons."

For inclusive branching fractions, e.g.,  $\Lambda_b \rightarrow \bar{\Lambda}_c \text{anything}$ , the values usually are multiplicities, not branching fractions. They can be greater than one.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0)$	$(5.8 \pm 0.8) \times 10^{-5}$	
$\Gamma_2$ $J/\psi(1S) \Lambda$		
$\Gamma_3$ $\psi(2S) \Lambda$		
$\Gamma_4$ $p D^0 \pi^-$	$(6.4 \pm 0.7) \times 10^{-4}$	
$\Gamma_5$ $p D^0 K^-$	$(4.7 \pm 0.8) \times 10^{-5}$	
$\Gamma_6$ $p J/\psi \pi^-$	$(2.6 \pm 0.5) \times 10^{-5}$	
$\Gamma_7$ $p J/\psi K^-$	$(3.2 \pm 0.6) \times 10^{-4}$	
$\Gamma_8$ $P_c(4380)^+ K^-$ , $P_c \rightarrow p J/\psi$ [a]	$(2.7 \pm 1.4) \times 10^{-5}$	
$\Gamma_9$ $P_c(4450)^+ K^-$ , $P_c \rightarrow p J/\psi$ [a]	$(1.3 \pm 0.4) \times 10^{-5}$	
$\Gamma_{10}$ $p \bar{K}^0 \pi^-$	$(1.3 \pm 0.4) \times 10^{-5}$	
$\Gamma_{11}$ $p K^0 K^-$	$< 3.5 \times 10^{-6}$	CL=90%
$\Gamma_{12}$ $\Lambda_c^+ \pi^-$	$(4.9 \pm 0.4) \times 10^{-3}$	S=1.2
$\Gamma_{13}$ $\Lambda_c^+ K^-$	$(3.59 \pm 0.30) \times 10^{-4}$	S=1.2
$\Gamma_{14}$ $\Lambda_c^+ a_1(1260)^-$	seen	
$\Gamma_{15}$ $\Lambda_c^+ D^-$	$(4.6 \pm 0.6) \times 10^{-4}$	
$\Gamma_{16}$ $\Lambda_c^+ D_s^-$	$(1.10 \pm 0.10) \%$	
$\Gamma_{17}$ $\Lambda_c^+ \pi^+ \pi^- \pi^-$	$(7.7 \pm 1.1) \times 10^{-3}$	S=1.1
$\Gamma_{18}$ $\Lambda_c(2595)^+ \pi^-$ , $\Lambda_c(2595)^+ \rightarrow \Lambda_c^+ \pi^+ \pi^-$	$(3.4 \pm 1.5) \times 10^{-4}$	
$\Gamma_{19}$ $\Lambda_c(2625)^+ \pi^-$ , $\Lambda_c(2625)^+ \rightarrow \Lambda_c^+ \pi^+ \pi^-$	$(3.3 \pm 1.3) \times 10^{-4}$	
$\Gamma_{20}$ $\Sigma_c(2455)^0 \pi^+ \pi^-$ , $\Sigma_c^0 \rightarrow \Lambda_c^+ \pi^-$	$(5.7 \pm 2.2) \times 10^{-4}$	
$\Gamma_{21}$ $\Sigma_c(2455)^{++} \pi^- \pi^-$ , $\Sigma_c^{++} \rightarrow \Lambda_c^+ \pi^+$	$(3.2 \pm 1.6) \times 10^{-4}$	
$\Gamma_{22}$ $\Lambda K^0 2\pi^+ 2\pi^-$		
$\Gamma_{23}$ $\Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything}$	[b] $(10.3 \pm 2.2) \%$	
$\Gamma_{24}$ $\Lambda_c^+ \ell^- \bar{\nu}_\ell$	$(6.2 \pm 1.4) \%$	
$\Gamma_{25}$ $\Lambda_c^+ \pi^+ \pi^- \ell^- \bar{\nu}_\ell$	$(5.6 \pm 3.1) \%$	
$\Gamma_{26}$ $\Lambda_c(2595)^+ \ell^- \bar{\nu}_\ell$	$(7.9 \pm 4.0) \times 10^{-3}$	
$\Gamma_{27}$ $\Lambda_c(2625)^+ \ell^- \bar{\nu}_\ell$	$(1.3 \pm 0.6) \%$	
$\Gamma_{28}$ $\Sigma_c(2455)^0 \pi^+ \ell^- \bar{\nu}_\ell$		
$\Gamma_{29}$ $\Sigma_c(2455)^{++} \pi^- \ell^- \bar{\nu}_\ell$		
$\Gamma_{30}$ $p h^-$	[c] $< 2.3 \times 10^{-5}$	CL=90%
$\Gamma_{31}$ $p \pi^-$	$(4.2 \pm 0.8) \times 10^{-6}$	
$\Gamma_{32}$ $p K^-$	$(5.1 \pm 1.0) \times 10^{-6}$	
$\Gamma_{33}$ $p D_s^-$	$< 4.8 \times 10^{-4}$	CL=90%
$\Gamma_{34}$ $p \mu^- \bar{\nu}_\mu$	$(4.1 \pm 1.0) \times 10^{-4}$	
$\Gamma_{35}$ $\Lambda \mu^+ \mu^-$	$(1.08 \pm 0.28) \times 10^{-6}$	
$\Gamma_{36}$ $\Lambda \gamma$	$< 1.3 \times 10^{-3}$	CL=90%
$\Gamma_{37}$ $\Lambda^0 \eta$	$(9 \pm 7) \times 10^{-6}$	
$\Gamma_{38}$ $\Lambda^0 \eta'(958)$	$< 3.1 \times 10^{-6}$	CL=90%

[a]  $P_c^+$  is a pentaquark-charmonium state.

[b] Not a pure measurement. See note at head of  $\Lambda_b^0$  Decay Modes.

[c] Here  $h^-$  means  $\pi^-$  or  $K^-$ .

## CONSTRAINED FIT INFORMATION

An overall fit to 10 branching ratios uses 12 measurements and one constraint to determine 7 parameters. The overall fit has a  $\chi^2 = 10.7$  for 6 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i/\Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_{13}$	94				
$x_{17}$	50	47			
$x_{24}$	14	14	7		
$x_{31}$	0	0	0	0	
$x_{32}$	0	0	0	0	83
	$x_{12}$	$x_{13}$	$x_{17}$	$x_{24}$	$x_{31}$

 $\Lambda_b^0$  BRANCHING RATIOS $\Gamma(J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0))/\Gamma_{\text{total}}$  $\Gamma_1/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>5.8 \pm 0.8</math> OUR AVERAGE</b>				
$6.01 \pm 0.60 \pm 0.58 \pm 0.28$		<sup>1</sup> ABAZOV	11o D0	$p \bar{p}$ at 1.96 TeV
$4.7 \pm 2.3 \pm 0.2$		<sup>2</sup> ABE	97b CDF	$p \bar{p}$ at 1.8 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$180 \pm 60 \pm 90$	16	ALBAJAR	91E UA1	$p \bar{p}$ at 630 GeV
<sup>1</sup> ABAZOV 11o uses $B(B^0 \rightarrow J/\psi K_S^0) \times B(b \rightarrow B^0) = (1.74 \pm 0.08) \times 10^{-4}$ to obtain the result. The $(\pm 0.08) \times 10^{-4}$ uncertainty of this product is listed as the last uncertainty of the measurement, $(\pm 0.28) \times 10^{-5}$ .				
<sup>2</sup> ABE 97b reports $[B(\Lambda_b^0 \rightarrow J/\psi \Lambda) \times B(b \rightarrow \Lambda_b^0)] / [B(B^0 \rightarrow J/\psi K_S^0) \times B(b \rightarrow B^0)] = 0.27 \pm 0.12 \pm 0.05$ . We multiply by our best value $B(B^0 \rightarrow J/\psi K_S^0) \times B(b \rightarrow B^0) = (1.74 \pm 0.08) \times 10^{-4}$ . Our first error is their experiment error and our second error is the systematic error from using our best value.				

 $\Gamma(\psi(2S) \Lambda)/\Gamma(J/\psi(1S) \Lambda)$  $\Gamma_3/\Gamma_2$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.50 \pm 0.03 \pm 0.02</math></b>	<sup>1</sup> AAD	15CH ATLS	$pp$ at 8 TeV
<sup>1</sup> AAD 15CH uses $B(J/\psi \rightarrow \mu^+ \mu^-) = (5.961 \pm 0.033) \times 10^{-2}$ (PDG 14). And $B(\psi(2S) \rightarrow \mu^+ \mu^-) = (7.89 \pm 0.17) \times 10^{-3}$ (PDG 14) is used assuming lepton universality.			

 $\Gamma(p D^0 \pi^-)/\Gamma_{\text{total}}$  $\Gamma_4/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
seen	52	BAR1	91 SFM	$D^0 \rightarrow K^- \pi^+$
seen		BASILE	81 SFM	$D^0 \rightarrow K^- \pi^+$

 $\Gamma(p D^0 K^-)/\Gamma(p D^0 \pi^-)$  $\Gamma_5/\Gamma_4$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>7.3 \pm 0.8 \pm 0.6</math></b>	AAIJ	14H LHCB	$pp$ at 7 TeV

 $\Gamma(p J/\psi \pi^-)/\Gamma(p J/\psi K^-)$  $\Gamma_6/\Gamma_7$ 

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>8.24 \pm 0.25 \pm 0.42</math></b>	AAIJ	14K LHCB	$pp$ at 7, 8 TeV

 $\Gamma(p J/\psi K^-)/\Gamma_{\text{total}}$  $\Gamma_7/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>3.17 \pm 0.04 \pm 0.57 \pm 0.45</math></b>	<sup>1</sup> AAIJ	16A LHCB	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 16A reported the measurement of $(3.17 \pm 0.04 \pm 0.07 \pm 0.34 \pm 0.45) \times 10^{-4}$ where the first uncertainty is statistical, the second is systematic, the third is due to the branching fraction of $B^0 \rightarrow J/\psi K^*(892)^0$ , and the fourth is due to the knowledge of $f_{\Lambda_b}/f_d$ . We combined in quadrature second to fourth uncertainties to a total systematic uncertainty.			

 $\Gamma(P_c(4380)^+ K^-)$  $\Gamma_8/\Gamma$ 

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.66 \pm 0.22 \pm 1.41 \pm 1.38</math></b>	<sup>1</sup> AAIJ	16A LHCB	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 16 total systematic includes the uncertainties on $f(P_c^+)$ and $B(\Lambda_b \rightarrow p J/\psi K^-)$ .			

 $\Gamma(P_c(4450)^+ K^-)$  $\Gamma_9/\Gamma$ 

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.30 \pm 0.16 \pm 0.42 \pm 0.39</math></b>	<sup>1</sup> AAIJ	16A LHCB	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 16 total systematic includes the uncertainties on $f(P_c^+)$ and $B(\Lambda_b \rightarrow p J/\psi K^-)$ .			

 $\Gamma(p \bar{K}^0 \pi^-)/\Gamma_{\text{total}}$  $\Gamma_{10}/\Gamma$ 

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.26 \pm 0.19 \pm 0.36</math></b>	<sup>1</sup> AAIJ	14Q LHCB	$pp$ at 7 TeV

<sup>1</sup> Used the normalizing mode branching fraction value of  $B(B^0 \rightarrow K^0 \pi^+ \pi^-) = (4.96 \pm 0.20) \times 10^{-5}$ .

 $\Gamma(p \bar{K}^0 K^-)/\Gamma_{\text{total}}$  $\Gamma_{11}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt; 3.5 \times 10^{-6}</math></b>	90	AAIJ	14Q LHCB	$pp$ at 7 TeV

# Baryon Particle Listings

$\Lambda_b^0$

$\Gamma(\Lambda_c^+ \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{12}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>4.9 ±0.4 OUR FIT</b>	Error includes scale factor of 1.2.					
<b>4.9 ±0.5 OUR AVERAGE</b>	Error includes scale factor of 1.5.					
$4.57^{+0.31}_{-0.30} \pm 0.23$	1	AAIJ	14I	LHCB	$p\bar{p}$ at 7 TeV	
$5.97 \pm 0.28 \pm 0.81$	2	AAIJ	14Q	LHCB	$p\bar{p}$ at 7 TeV	
$8.8 \pm 2.8 \pm 1.5$	3	ABULENCIA	07B	CDF	$p\bar{p}$ at 1.96 TeV	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
seen	3	ABREU	96N	DLPH	$\Lambda_c^+ \rightarrow p K^- \pi^+$	
seen	4	BUSKULIC	96L	ALEP	$\Lambda_c^+ \rightarrow p K^- \pi^+,$ $\rho^0 \pi^-, \Lambda \pi^+ \pi^-$	

<sup>1</sup> AAIJ 14I reports  $(4.30 \pm 0.03^{+0.12}_{-0.11} \pm 0.26 \pm 0.21) \times 10^{-3}$  from a measurement of  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-)/\Gamma_{\text{total}}] \times [\text{B}(\bar{B}^0 \rightarrow D^- \pi^+)]$  assuming  $\text{B}(\bar{B}^0 \rightarrow D^- \pi^+) = (2.68 \pm 0.13) \times 10^{-3}$ , which we rescale to our best value  $\text{B}(\bar{B}^0 \rightarrow D^- \pi^+) = (2.52 \pm 0.13) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Uses information on  $f_{\text{baryon}}/f_d$  from measurement in semileptonic decays by the same authors.

<sup>2</sup> Obtained using the branching fraction of  $\Lambda_c^+ \rightarrow p K^- \pi^+$  decay.

<sup>3</sup> The result is obtained from  $(f_{\text{baryon}}/f_d) (\text{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-)/\text{B}(\bar{B}^0 \rightarrow D^+ \pi^-)) = 0.82 \pm 0.08 \pm 0.11 \pm 0.22$ , assuming  $f_{\text{baryon}}/f_d = 0.25 \pm 0.04$  and  $\text{B}(\bar{B}^0 \rightarrow D^+ \pi^-) = (2.68 \pm 0.13) \times 10^{-3}$ .

$\Gamma(pD^0\pi^-)/\Gamma(\Lambda_c^+\pi^-)$					$\Gamma_4/\Gamma_{12}$	
VALUE	DOCUMENT ID		TECN	COMMENT		
<b>0.130±0.007±0.007</b>	<sup>1</sup>	AAIJ	14H	LHCB	$p\bar{p}$ at 7 TeV	
<sup>1</sup> AAIJ 14H reports $[\Gamma(\Lambda_b^0 \rightarrow pD^0\pi^-)/\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+\pi^-)] \times [\text{B}(D^0 \rightarrow K^-\pi^+)] / [\text{B}(\Lambda_c^+ \rightarrow pK^-\pi^+)] = (8.06 \pm 0.23 \pm 0.35) \times 10^{-2}$ which we multiply or divide by our best values $\text{B}(D^0 \rightarrow K^-\pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ , $\text{B}(\Lambda_c^+ \rightarrow pK^-\pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.						

$\Gamma(\Lambda_c^+ K^-)/\Gamma_{\text{total}}$					$\Gamma_{13}/\Gamma$	
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT			
<b>3.59 ± 0.30 OUR FIT</b>	Error includes scale factor of 1.2.					
<b>3.55 ± 0.44 ± 0.50</b>	<sup>1</sup> AAJ1	14Q	LHCB	$p\bar{p}$ at 7 TeV		
<sup>1</sup> Obtained using the branching fraction of $\Lambda_c^+ \rightarrow p K^- \pi^+$ decay.						

$\Gamma(\Lambda_c^+ K^-)/\Gamma(\Lambda_c^+ \pi^-)$	$\Gamma_{13}/\Gamma_{12}$			
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT	
<b>7.31±0.22 OUR FIT</b>				
<b>7.31±0.16±0.16</b>	AAIJ	14H	LHCB	$p\bar{p}$ at 7 TeV

$\Gamma(\Lambda_c^+ a_1(1260)^-)/\Gamma_{\text{total}}$					$\Gamma_{14}/\Gamma$	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
seen	1	ABREU	96N	DLPH	$\Lambda_c^+ \rightarrow p K^- \pi^+, a_1^- \rightarrow \rho^0 \pi^- \rightarrow \pi^+ \pi^- \pi^-$	

$\Gamma(\Lambda_c^+ D_s^-)/\Gamma_{\text{total}}$	$\Gamma_{16}/\Gamma$		
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.1 <math>\pm</math> 0.1</b>	<sup>1</sup> AAIJ	14AA	LHCB $p\bar{p}$ at 7 TeV
<sup>1</sup> Uses $\text{B}(\bar{B}^0 \rightarrow D^+ D_s^-) = (7.2 \pm 0.8) \times 10^{-3}$ and their measured $\text{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-)/\text{B}(\bar{B}^0 \rightarrow D^+ \pi^-)$ values.			

$\Gamma(\Lambda_c^+ D^-)/\Gamma(\Lambda_c^+ D_s^-)$					$\Gamma_{15}/\Gamma_{16}$	
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>0.042 ±0.003 ±0.003</b>	AAIJ	14AA	LHCB	$p\bar{p}$ at 7 TeV		

$\Gamma(\Lambda_c^+ \pi^+ \pi^- \pi^-)/\Gamma_{\text{total}}$					$\Gamma_{17}/\Gamma$	
VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>7.7<math>\pm</math>1.1 OUR FIT</b>	Error includes scale factor of 1.1.					
<b>14.9<math>\pm</math>3.8 3.2<math>\pm</math>1.2</b>	1	AALTONEN	12A	CDF	$p\bar{p}$ at 1.96 TeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
seen	90	BARI	91	SFM	$\Lambda_c^+ \rightarrow p K^- \pi^+$	

<sup>1</sup> AALTONEN 12A reports  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^+ \pi^- \pi^-)/\Gamma_{\text{total}}] / [\text{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-)] = 3.04 \pm 0.33^{+0.70}_{-0.55}$  which we multiply by our best value  $\text{B}(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-) = (4.9 \pm 0.4) \times 10^{-3}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Lambda_c^+ \pi^+ \pi^- \pi^-)/\Gamma(\Lambda_c^+ \pi^-)$				$\Gamma_{17}/\Gamma_{12}$	
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>1.56±0.21 OUR FIT</b>					
<b>1.43±0.16±0.13</b>	AAIJ	11E	LHCB	$p\bar{p}$ at 7 TeV	

$\Gamma(\Lambda_c(2595)^+ \pi^-, \Lambda_c(2595)^+ \rightarrow \Lambda_c^+ \pi^+ \pi^-)/\Gamma(\Lambda_c^+ \pi^+ \pi^- \pi^-)$					$\Gamma_{18}/\Gamma_{17}$	
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT			
<b>4.4 ±1.7 ±<math>\frac{0.6}{0.4}</math></b>	AAIJ	11E	LHCB	$p\bar{p}$ at 7 TeV		

$\Gamma(\Lambda_c(2625)^+ \pi^-, \Lambda_c(2625)^+ \rightarrow \Lambda_c^+ \pi^+ \pi^-)/\Gamma(\Lambda_c^+ \pi^+ \pi^- \pi^-)$					$\Gamma_{19}/\Gamma_{17}$	
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT			
<b>4.3 ±1.5 ±0.4</b>	AAIJ	11E	LHCB	$p\bar{p}$ at 7 TeV		

$\Gamma(\Sigma_c(2455)^0 \pi^+ \pi^-, \Sigma_c^0 \rightarrow \Lambda_c^+ \pi^-)/\Gamma(\Lambda_c^+ \pi^+ \pi^- \pi^-)$					$\Gamma_{20}/\Gamma_{17}$	
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT			
<b>7.4 ±2.4 ±1.2</b>	AAIJ	11E	LHCB	$p\bar{p}$ at 7 TeV		

$\Gamma(\Sigma_c(2455)^{++} \pi^- \pi^-, \Sigma_c^{++} \rightarrow \Lambda_c^+ \pi^+)/\Gamma(\Lambda_c^+ \pi^+ \pi^- \pi^-)$					$\Gamma_{21}/\Gamma_{17}$	
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT			
<b>4.2 ±1.8 ±0.7</b>	AAIJ	11E	LHCB	$p\bar{p}$ at 7 TeV		

$\Gamma(\Lambda K^0 2\pi^+ 2\pi^-)/\Gamma_{\text{total}}$					$\Gamma_{22}/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
seen	4	<sup>1</sup> ARENTON	86	FMPS	$\Lambda K^0_S 2\pi^+ 2\pi^-$

<sup>1</sup> See the footnote to the ARENTON 86 mass value.

$\Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})/\Gamma_{\text{total}}$					$\Gamma_{23}/\Gamma$	
The values and averages in this section serve only to show what values result if one assumes our $B(b \rightarrow b\text{-baryon})$ . They cannot be thought of as measurements since the underlying product branching fractions were also used to determine $B(b \rightarrow b\text{-baryon})$ as described in the note on "Production and Decay of $b$ -Flavored Hadrons."						
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT		
<b>0.103±0.022 OUR AVERAGE</b>						
0.097±0.018±0.014		<sup>1</sup> BARATE	98D	ALEP	$e^+ e^- \rightarrow Z$	
0.13 $\begin{smallmatrix} +0.05 \\ -0.04 \end{smallmatrix}$ ±0.02	29	<sup>2</sup> ABREU	95s	DLPH	$e^+ e^- \rightarrow Z$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
0.085±0.021±0.012	55	<sup>3</sup> BUSKULIC	95L	ALEP	Repl. by BARATE 98D	
0.17 ±0.06 ±0.02	21	<sup>4</sup> BUSKULIC	92E	ALEP	$\Lambda_c^+ \rightarrow p K^- \pi^+$	

<sup>1</sup> BARATE 98D reports  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.0086 \pm 0.0007 \pm 0.0014$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Measured using  $\Lambda_c \ell^-$  and  $\Lambda \ell^+ \ell^-$ .

<sup>2</sup> ABREU 95S reports  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.0118 \pm 0.0026^{+0.0031}_{-0.0021}$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>3</sup> BUSKULIC 95L reports  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.00755 \pm 0.0014 \pm 0.0012$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>4</sup> BUSKULIC 92E reports  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell \text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.015 \pm 0.0035 \pm 0.0045$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Superseded by BUSKULIC 95L.

$\Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)/\Gamma_{\text{total}}$					$\Gamma_{24}/\Gamma$	
VALUE	DOCUMENT ID	TECN	COMMENT			
$0.062^{+0.014}_{-0.013}$ OUR FIT						
$0.050^{+0.011}_{-0.008} \pm^{+0.016}_{-0.012}$	<sup>1</sup> ABDALLAH	04A	DLPH	$e^+e^- \rightarrow Z^0$		
<sup>1</sup> Derived from a combined likelihood and event rate fit to the distribution of the $l_{\text{sgur}}$ -Wise variable and using HQET. The slope of the form factor is measured to be $\rho^2 = 2.03 \pm 0.46^{+0.72}_{-1.00}$ .						

$\Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)/\Gamma(\Lambda_c^+ \pi^-)$	$\Gamma_{24}/\Gamma_{12}$				
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>12.7<math>^{+3.1}_{-2.7}</math> OUR FIT</b>					
<b>16.6<math>\pm 3.0^{+2.8}_{-3.6}</math></b>	AALTONEN	09E	CDF	$p\bar{p}$ at 1.96 TeV	

$\Gamma(\Lambda_c^+ \pi^+ \pi^- \ell^- \bar{\nu}_\ell) / \Gamma_{\text{total}}$	$\Gamma_{25} / \Gamma$			
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.056^{+0.031}_{-0.030}$	<sup>1</sup> ABDALLAH	04A	DLPH	$e^+ e^- \rightarrow Z^0$
<sup>1</sup> Derived from the fraction of $\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell) / (\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \ell^- \bar{\nu}_\ell) + \Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^+ \pi^- \ell^- \bar{\nu}_\ell)) = 0.47^{+0.10+0.07}_{-0.08-0.06}$ .				

$\Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)/[\Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell) + \Gamma(\Lambda_c^+ \pi^+ \pi^- \ell^- \bar{\nu}_\ell)]$					$\Gamma_{24}/(\Gamma_{24} + \Gamma_{25})$	
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>0.47 ±<math>\frac{0.10}{0.08}</math> ±<math>\frac{0.07}{0.06}</math></b>	ABDALLAH	04A	DLPH	$e^+ e^- \rightarrow Z^0$		

$\Gamma(\Lambda_c(2595)^+ \ell^- \bar{\nu}_\ell)/\Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)$					$\Gamma_{26}/\Gamma_{24}$	
VALUE	DOCUMENT ID	TECN	COMMENT			
<b>0.126 ±0.033 ±<math>\frac{0.047}{0.038}</math></b>	AALTONEN	09E	CDF	$p\bar{p}$ at 1.96 TeV		

See key on page 601

## Baryon Particle Listings

 $\Lambda_b^0$ 

$\Gamma(\Lambda_c(2625)^+ \ell^- \bar{\nu}_\ell) / \Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)$	$\Gamma_{27} / \Gamma_{24}$
VALUE	DOCUMENT ID
$0.210 \pm 0.042 \pm 0.071$ $0.050$	AALTONEN 09E CDF $p\bar{p}$ at 1.96 TeV

$[\frac{1}{2}\Gamma(\Sigma_c(2455)^0 \pi^+ \ell^- \bar{\nu}_\ell) + \frac{1}{2}\Gamma(\Sigma_c(2455)^{++} \pi^- \ell^- \bar{\nu}_\ell)] / \Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)$ $(\frac{1}{2}\Gamma_{28} + \frac{1}{2}\Gamma_{29}) / \Gamma_{24}$	
VALUE	DOCUMENT ID
$0.054 \pm 0.022 \pm 0.021$ $-0.018$	AALTONEN 09E CDF $p\bar{p}$ at 1.96 TeV

$\Gamma(p h^-) / \Gamma_{\text{total}}$	$\Gamma_{30} / \Gamma$
VALUE	CL%
$< 2.3 \times 10^{-5}$	90
	<sup>1</sup> ACOSTA 05o CDF $p\bar{p}$ at 1.96 TeV
<sup>1</sup> Assumes $f_\Lambda / f_d = 0.25$ , and equal momentum distribution for $\Lambda_b$ and $B$ mesons.	

$\Gamma(p\pi^-) / \Gamma_{\text{total}}$	$\Gamma_{31} / \Gamma$
VALUE (units $10^{-6}$ )	CL%
<b><math>4.2 \pm 0.8</math> OUR FIT</b> <b><math>3.7 \pm 0.8 \pm 0.6</math></b>	
	<sup>1</sup> AALTONEN 09c CDF $p\bar{p}$ at 1.96 TeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 50$  90 <sup>2</sup> BUSKULIC 96v ALEP  $e^+ e^- \rightarrow Z$

<sup>1</sup> AALTONEN 09c reports  $[\Gamma(\Lambda_b^0 \rightarrow p\pi^-) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow K^+ \pi^-)] \times [B(\bar{B} \rightarrow b\text{-baryon})] / [B(\bar{B} \rightarrow B^0)] = 0.042 \pm 0.007 \pm 0.006$  which we multiply or divide by our best values  $B(B^0 \rightarrow K^+ \pi^-) = (1.96 \pm 0.05) \times 10^{-5}$ ,  $B(\bar{B} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ ,  $B(\bar{B} \rightarrow B^0) = (40.4 \pm 0.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

<sup>2</sup> BUSKULIC 96v assumes PDG 96 production fractions for  $B^0$ ,  $B^+$ ,  $B_s$ ,  $b$  baryons.

$\Gamma(pK^-) / \Gamma_{\text{total}}$	$\Gamma_{32} / \Gamma$
VALUE (units $10^{-6}$ )	CL%
<b><math>5.1 \pm 1.0</math> OUR FIT</b> <b><math>5.9 \pm 1.1 \pm 0.9</math></b>	
	<sup>1</sup> AALTONEN 09c CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$< 360$ 90 <sup>2</sup> ADAM 96d DLPH $e^+ e^- \rightarrow Z$	
$< 50$ 90 <sup>3</sup> BUSKULIC 96v ALEP $e^+ e^- \rightarrow Z$	
<sup>1</sup> AALTONEN 09c reports $[\Gamma(\Lambda_b^0 \rightarrow pK^-) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow K^+ \pi^-)] \times [B(\bar{B} \rightarrow b\text{-baryon})] / [B(\bar{B} \rightarrow B^0)] = 0.066 \pm 0.009 \pm 0.008$ which we multiply or divide by our best values $B(B^0 \rightarrow K^+ \pi^-) = (1.96 \pm 0.05) \times 10^{-5}$ , $B(\bar{B} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ , $B(\bar{B} \rightarrow B^0) = (40.4 \pm 0.6) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.	
<sup>2</sup> ADAM 96d assumes $f_{B^0} = f_{B^-} = 0.39$ and $f_{B_s} = 0.12$ .	
<sup>3</sup> BUSKULIC 96v assumes PDG 96 production fractions for $B^0$ , $B^+$ , $B_s$ , $b$ baryons.	

$\Gamma(p\pi^-) / \Gamma(pK^-)$	$\Gamma_{31} / \Gamma_{32}$
VALUE	DOCUMENT ID
<b><math>0.84 \pm 0.09</math> OUR FIT</b> <b><math>0.86 \pm 0.08 \pm 0.05</math></b>	AAIJ 12AR LHCB $pp$ at 7 TeV

$\Gamma(pD_s^-) / \Gamma_{\text{total}}$	$\Gamma_{33} / \Gamma$
VALUE	CL%
$< 4.8 \times 10^{-4}$	90
	AAIJ 14q LHCB $pp$ at 7 TeV

$\Gamma(p\mu^- \bar{\nu}_\mu) / \Gamma_{\text{total}}$	$\Gamma_{34} / \Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID
<b><math>4.1 \pm 1.0</math></b>	<sup>1</sup> AAIJ 15BG LHCB $pp$ at 8 TeV

<sup>1</sup> The ratio of  $B(\Lambda_b^0 \rightarrow p\mu^- \bar{\nu}_\mu)$  to  $B(\Lambda_b^0 \rightarrow \Lambda_c^+ \mu^- \bar{\nu}_\mu)$  is measured within a restricted  $q^2$  region. Combined with theoretical calculations of the form factors and the previously measured value of  $|V_{cb}|$ , the first  $|V_{ub}| = (3.27 \pm 0.15 \pm 0.16 \pm 0.06) \times 10^{-3}$  measurement from the  $\Lambda_b$  decay is obtained, consistent with the exclusively measured world averages.

$\Gamma(p\mu^- \bar{\nu}_\mu) / \Gamma(\Lambda_c^+ \ell^- \bar{\nu}_\ell)$	$\Gamma_{34} / \Gamma_{24}$
VALUE (units $10^{-2}$ )	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$1.0 \pm 0.04 \pm 0.08$	<sup>1</sup> AAIJ 15BG LHCB $pp$ at 8 TeV

<sup>1</sup> This measurement is a ratio of  $\Gamma(\Lambda_b^0 \rightarrow p\mu^- \bar{\nu}_\mu)[q^2 > 15 \text{ GeV}/c^2]$  to  $\Gamma(\Lambda_b^0 \rightarrow \Lambda_c^+ \mu^- \bar{\nu}_\mu)[q^2 > 7 \text{ GeV}/c^2]$  within a restricted  $q^2$  region. Combined with theoretical calculations of the form factors and the previously measured value of  $|V_{cb}|$ , the first  $|V_{ub}| = (3.27 \pm 0.15 \pm 0.16 \pm 0.06) \times 10^{-3}$  measurement from the  $\Lambda_b$  decay is obtained, consistent with the exclusively measured world averages.

$\Gamma(\Lambda\mu^+ \mu^-) / \Gamma_{\text{total}}$	$\Gamma_{35} / \Gamma$
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>10.8 \pm 2.8</math> OUR AVERAGE</b>	
$9.6 \pm 1.6 \pm 2.5$	<sup>1</sup> AAIJ 13AJ LHCB $pp$ at 7 TeV
$17.3 \pm 4.2 \pm 5.5$	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV

<sup>1</sup> Uses  $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ . This measurement comes from the sum of the differential rates in  $q^2$  regions excluding those corresponding to  $J/\psi$  and  $\psi(2S)$  ([8.68, 10.09] and [12.86, 14.18]  $\text{GeV}^2/c^4$ ).

$\Gamma(\Lambda\gamma) / \Gamma_{\text{total}}$	$\Gamma_{36} / \Gamma$
VALUE	CL%
$< 1.3 \times 10^{-3}$	90
	ACOSTA 02G CDF $p\bar{p}$ at 1.8 TeV

$\Gamma(\Lambda^0 \eta) / \Gamma_{\text{total}}$	$\Gamma_{37} / \Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID
<b><math>9 \pm 7 \pm 1</math></b>	<sup>1</sup> AAIJ 15AH LHCB $pp$ at 7, 8 TeV

<sup>1</sup> AAIJ 15AH reports  $[\Gamma(\Lambda_b^0 \rightarrow \Lambda^0 \eta) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow \eta' K^0)] = 0.142 \pm 0.11$  which we multiply by our best value  $B(B^0 \rightarrow \eta' K^0) = (6.6 \pm 0.4) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. The single uncertainty quoted with the original measurement combines in quadrature statistical and systematic uncertainties.

$\Gamma(\Lambda^0 \eta'(958)) / \Gamma_{\text{total}}$	$\Gamma_{38} / \Gamma$
VALUE	CL%
$< 3.1 \times 10^{-6}$	90
	<sup>1</sup> AAIJ 15AH LHCB $pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AH reports $[\Gamma(\Lambda_b^0 \rightarrow \Lambda^0 \eta'(958)) / \Gamma_{\text{total}}] / [B(B^0 \rightarrow \eta' K^0)] < 0.047$ which we multiply by our best value $B(B^0 \rightarrow \eta' K^0) = 6.6 \times 10^{-5}$ .	

PARTIAL BRANCHING FRACTIONS IN  $\Lambda_b \rightarrow \Lambda\mu^+ \mu^-$ 

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (q^2 < 2.0 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>0.71 \pm 0.27</math> OUR AVERAGE</b>	
$0.72 \pm 0.24 \pm 0.14$	<sup>1</sup> AAIJ 15AE LHCB $pp$ at 7, 8 TeV
$0.15 \pm 2.01 \pm 0.05$	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.56 \pm 0.76 \pm 0.80$	<sup>2</sup> AAIJ 13AJ LHCB Repl. by AAIJ 15AE
<sup>1</sup> AAIJ 15AE measurement covers $0.1 < q^2 < 2.0 \text{ GeV}^2/c^4$ .	
<sup>2</sup> Uses $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ .	

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (2.0 < q^2 < 4.3 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>0.28 \pm 0.28</math> OUR AVERAGE</b> <b><math>-0.21</math></b>	
$0.253 \pm 0.276 \pm 0.046$	<sup>1</sup> AAIJ 15AE LHCB $pp$ at 7, 8 TeV
$1.8 \pm 1.7 \pm 0.6$	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$0.71 \pm 0.60 \pm 0.23$	<sup>2</sup> AAIJ 13AJ LHCB Repl. by AAIJ 15AE
<sup>1</sup> AAIJ 15AE measurement covers $2.0 < q^2 < 4.0 \text{ GeV}^2/c^4$ .	
<sup>2</sup> Uses $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ .	

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (q^2 < 4.3 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>2.7 \pm 2.5 \pm 0.9</math></b>	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (4.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>0.04 \pm 0.18</math> OUR AVERAGE</b> <b><math>-0.00</math></b>	AAIJ 15AE LHCB $pp$ at 7, 8 TeV

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (1.0 < q^2 < 6.0 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>0.47 \pm 0.31</math> OUR AVERAGE</b> <b><math>-0.27</math></b>	
$0.45 \pm 0.30 \pm 0.10$	<sup>1</sup> AAIJ 15AE LHCB $pp$ at 7 and 8 TeV
$1.3 \pm 2.1 \pm 0.4$	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV
<sup>1</sup> AAIJ 15AE measurement covers $1.1 < q^2 < 6.0 \text{ GeV}^2/c^4$ .	

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (6.0 < q^2 < 8.0 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>0.50 \pm 0.24</math> OUR AVERAGE</b> <b><math>-0.22</math></b>	AAIJ 15AE LHCB $pp$ at 7, 8 TeV

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (4.3 < q^2 < 8.68 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>0.5 \pm 0.7</math> OUR AVERAGE</b>	
$0.66 \pm 0.74 \pm 0.18$	<sup>1</sup> AAIJ 13AJ LHCB $pp$ at 7 TeV
$-0.2 \pm 1.6 \pm 0.1$	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV
<sup>1</sup> Uses $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ .	

$B(\Lambda_b \rightarrow \Lambda\mu^+ \mu^-) (10.09 < q^2 < 12.86 \text{ GeV}^2/c^4)$	
VALUE (units $10^{-7}$ )	DOCUMENT ID
<b><math>2.2 \pm 0.6</math> OUR AVERAGE</b>	
$2.08 \pm 0.42 \pm 0.42$	<sup>1</sup> AAIJ 15AE LHCB $pp$ at 7, 8 TeV
$3.0 \pm 1.5 \pm 1.0$	AALTONEN 11AI CDF $p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$1.55 \pm 0.58 \pm 0.55$	<sup>2</sup> AAIJ 13AJ LHCB Repl. by AAIJ 15AE
<sup>1</sup> AAIJ 15AE measurement covers $11.0 < q^2 < 12.5 \text{ GeV}^2/c^4$ .	
<sup>2</sup> Uses $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ .	

# Baryon Particle Listings

## $\Lambda_b^0, \Lambda_b(5912)^0$

### $B(\Lambda_b \rightarrow \Lambda \mu^+ \mu^-)$ ( $14.18 < q^2 < 16.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>1.7 \pm 0.5</math> OUR AVERAGE</b>			Error includes scale factor of 1.1.
$2.04^{+0.35}_{-0.33} \pm 0.42$	<sup>1</sup> AAIJ	15AE LHCb	$pp$ at 7, 8 TeV
$1.0 \pm 0.7 \pm 0.3$	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$1.44 \pm 0.44 \pm 0.42$	<sup>2</sup> AAIJ	13AJ LHCb	Repl. by AAIJ 15AE
<sup>1</sup> AAIJ 15AE measurement covers $15.0 < q^2 < 16.0 \text{ GeV}^2/c^4$ .			
<sup>2</sup> Uses $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ .			

### $B(\Lambda_b \rightarrow \Lambda \mu^+ \mu^-)$ ( $16.0 < q^2 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>7.0 \pm 1.9 \pm 2.2</math></b>	AALTONEN	11AI CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$4.73 \pm 0.77 \pm 1.25$	<sup>1,2</sup> AAIJ	13AJ LHCb	Repl. by AAIJ 15AE
<sup>1</sup> Uses $B(\Lambda_b^0 \rightarrow J/\psi \Lambda) = (6.2 \pm 1.4) \times 10^{-4}$ .			
<sup>2</sup> Requires $16.00 < q^2 < 20.30 \text{ GeV}^2/c^4$ .			

### $B(\Lambda_b \rightarrow \Lambda \mu^+ \mu^-)$ ( $18.0 < q^2 < 20.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>2.44 \pm 0.28 \pm 0.50</math></b>	AAIJ	15AE LHCb	$pp$ at 7, 8 TeV

### $B(\Lambda_b \rightarrow \Lambda \mu^+ \mu^-)$ ( $15.0 < q^2 < 20.0 \text{ GeV}^2/c^4$ )

VALUE (units $10^{-7}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>6.00 \pm 0.45 \pm 1.25</math></b>	AAIJ	15AE LHCb	$pp$ at 7, 8 TeV

## CP VIOLATION

$A_{CP}$  is defined as

$$A_{CP} = \frac{B(\Lambda_b^0 \rightarrow f) - B(\bar{\Lambda}_b^0 \rightarrow \bar{f})}{B(\Lambda_b^0 \rightarrow f) + B(\bar{\Lambda}_b^0 \rightarrow \bar{f})},$$

the  $CP$ -violation asymmetry of exclusive  $\Lambda_b^0$  and  $\bar{\Lambda}_b^0$  decay.

### $A_{CP}(\Lambda_b \rightarrow \rho \pi^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.06 \pm 0.07</math> OUR AVERAGE</b>			
$0.06 \pm 0.07 \pm 0.03$	AALTONEN	14P CDF	$p\bar{p}$ at 1.96 TeV
$0.03 \pm 0.17 \pm 0.05$	AALTONEN	11N CDF	$p\bar{p}$ at 1.96 TeV

### $A_{CP}(\Lambda_b \rightarrow \rho K^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.00 \pm 0.19</math> OUR AVERAGE</b>			Error includes scale factor of 2.4.
$-0.10 \pm 0.08 \pm 0.04$	AALTONEN	14P CDF	$p\bar{p}$ at 1.96 TeV
$0.37 \pm 0.17 \pm 0.03$	AALTONEN	11N CDF	$p\bar{p}$ at 1.96 TeV

### $A_{CP}(\Lambda_b \rightarrow \rho \bar{K}^0 \pi^-)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.22 \pm 0.13 \pm 0.03</math></b>	AAIJ	14Q LHCb	$pp$ at 7 TeV

### $\Delta A_{CP}(J/\psi \rho \pi^- / K^-) \equiv A_{CP}(J/\psi \rho \pi^-) - A_{CP}(J/\psi \rho K^-)$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>5.7 \pm 2.4 \pm 1.2</math></b>	AAIJ	14K LHCb	$pp$ at 7, 8 TeV

## $\Lambda_b^0$ DECAY PARAMETERS

See the note on "Baryon Decay Parameters" in the neutron Listings.

### $\alpha$ decay parameter for $\Lambda_b \rightarrow J/\psi \Lambda$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.18 \pm 0.13</math> OUR AVERAGE</b>			
$0.30 \pm 0.16 \pm 0.06$	<sup>1</sup> AAD	14L ATLS	$pp$ at 7 TeV
$0.05 \pm 0.17 \pm 0.07$	<sup>2</sup> AAIJ	13AG LHCb	$pp$ at 7 TeV

<sup>1</sup> An angular analysis of  $\Lambda_b \rightarrow J/\psi \Lambda$  decay is performed and magnitudes of all helicity amplitudes are also reported.

<sup>2</sup> An angular analysis of  $\Lambda_b \rightarrow J/\psi \Lambda$  decay is performed and a  $\Lambda_b$  transverse production polarization of  $0.06 \pm 0.07 \pm 0.02$  is also reported.

### $A_{FB}^e(\mu\mu)$ in $\Lambda_b \rightarrow \Lambda \mu^+ \mu^-$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.05 \pm 0.09 \pm 0.03</math></b>	<sup>1</sup> AAIJ	15AE LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AE measurement covers $15.0 < q^2 < 20.0 \text{ GeV}^2/c^4$ .			

### $A_{FB}^h(\rho\pi)$ in $\Lambda_b \rightarrow \Lambda(\rho\pi) \mu^+ \mu^-$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.29 \pm 0.07 \pm 0.03</math></b>	<sup>1</sup> AAIJ	15AE LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AE measurement covers $15.0 < q^2 < 20.0 \text{ GeV}^2/c^4$ .			

### $f_L(\mu\mu)$ longitudinal polarization fraction in $\Lambda_b \rightarrow \Lambda \mu^+ \mu^-$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.61^{+0.11}_{-0.14} \pm 0.03</math></b>	<sup>1</sup> AAIJ	15AE LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> AAIJ 15AE measurement covers $15.0 < q^2 < 20.0 \text{ GeV}^2/c^4$ .			

## FORWARD-BACKWARD ASYMMETRIES

The forward-backward assymetry is defined as  $A_{FB}(\Lambda_b^0) = [N(F) - N(B)] / [N(F) + N(B)]$ , where the forward (F) direction corresponds to a particle ( $\Lambda_b^0$  or  $\Lambda_b^-$ ) sharing valence quark flavors with a beam particle with the same sign of rapidity.

### $A_{FB}(\Lambda_b^0 \rightarrow J/\psi \Lambda)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.04 \pm 0.07 \pm 0.02</math></b>	<sup>1</sup> ABAZOV	15I D0	$pp$ at 1.96 TeV
<sup>1</sup> The measured asymmetry integrated over rapidity y in the range of $0.1 <  y  < 2.0$ .			

## $\Lambda_b^0$ REFERENCES

AAIJ	16	JHEP 1601 012	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	16A	CPC 40 011001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAD	15CH	PL B751 63	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAIJ	15AE	JHEP 1506 115	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15AH	JHEP 1509 006	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	15BG	NATP 11 743	R. Aaij <i>et al.</i>	(LHCb Collab.)
ABAZOV	15I	PR D91 072008	V.M. Abazov <i>et al.</i>	(D0 Collab.)
AAD	14L	PR D89 092009	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAIJ	14AA	PRL 112 202001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14E	JHEP 1404 114	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14H	PR D89 032001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14I	JHEP 1408 143	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14K	JHEP 1407 103	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14Q	JHEP 1404 087	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	14U	PL B734 122	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	14B	PR D89 072014	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	14P	PRL 113 242001	T. Aaltonen <i>et al.</i>	(CDF Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)
AAD	13U	PR D87 032002	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAIJ	13AG	PL B724 27	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13AJ	PL B725 25	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13AV	PRL 110 182001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	13BB	PRL 111 102003	R. Aaij <i>et al.</i>	(LHCb Collab.)
CHATRCHYAN	13AC	JHEP 1307 163	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
AAIJ	12AR	JHEP 1210 037	R. Aaij <i>et al.</i>	(LHCb Collab.)
AAIJ	12E	PL B708 241	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	12A	PR D85 032003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	12U	PR D85 112003	V.M. Abazov <i>et al.</i>	(D0 Collab.)
AAIJ	11E	PR D84 092001	R. Aaij <i>et al.</i>	(LHCb Collab.)
Also		PR D85 039904 (errat.)	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	11	PRL 106 121804	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11AI	PRL 107 201802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11N	PRL 106 181802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	11O	PR D84 031102	V.M. Abazov <i>et al.</i>	(D0 Collab.)
AALTONEN	10B	PRL 104 102002	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	09C	PRL 103 031801	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	09E	PR D79 032001	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	07S	PRL 99 142001	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	07U	PRL 99 182001	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABULENCIA	07A	PRL 98 122001	A. Abulencia <i>et al.</i>	(FNAL CDF Collab.)
ABULENCIA	07B	PRL 98 122002	A. Abulencia <i>et al.</i>	(FNAL CDF Collab.)
ACOSTA	06	PRL 96 202001	D. Acosta <i>et al.</i>	(CDF Collab.)
ABAZOV	05C	PRL 94 102001	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ACOSTA	05O	PR D72 051104	D. Acosta <i>et al.</i>	(CDF Collab.)
ABDALLAH	04A	PL B585 63	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ACOSTA	02G	PR D66 112002	D. Acosta <i>et al.</i>	(CDF Collab.)
ABREU	99W	EPJ C10 185	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	98G	PL B426 161	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BARATE	98D	EPJ C2 197	R. Barate <i>et al.</i>	(ALEPH Collab.)
ABE	97B	PR D55 1142	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	96M	PRL 77 1439	F. Abe <i>et al.</i>	(CDF Collab.)
ABREU	96D	ZPHY C71 199	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	96N	PL B374 351	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ADAM	96D	ZPHY C72 207	W. Adam <i>et al.</i>	(DELPHI Collab.)
BUSKULIC	96L	PL B380 442	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	96V	PL B384 471	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
PDG	96	PR D54 1	R. M. Barnett <i>et al.</i>	(PDG Collab.)
ABREU	95K	ZPHY C68 375	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	95K	PL B353 402	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	95L	PL B357 685	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ABE	93B	PR D47 R2639	F. Abe <i>et al.</i>	(CDF Collab.)
BUSKULIC	92E	PL B294 145	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ALBAJAR	91E	PL B273 540	C. Albajar <i>et al.</i>	(UA1 Collab.)
BARI	91	NC 104A 1787	G. Bari <i>et al.</i>	(CERN R422 Collab.)
ARENTON	86	NP B274 707	M.W. Arenton <i>et al.</i>	(ARIZ, NDAM, VAND)
BASILE	81	LNC 31 97	M. Basile <i>et al.</i>	(CERN R415 Collab.)

## $\Lambda_b(5912)^0$

$$J^P = \frac{1}{2}^-$$

Status: \*\*\*

Quantum numbers are based on quark model expectations.

## $\Lambda_b(5912)^0$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>5912.11 \pm 0.13 \pm 0.23</math></b>	<sup>1,2</sup> AAIJ	12AL LHCb	$pp$ at 7 TeV

<sup>1</sup> Observed in  $\Lambda_b(5912)^0 \rightarrow \Lambda_b^0 \pi^+ \pi^-$  decays with  $17.6 \pm 4.8$  candidates with a significance of 5.2 sigma.

<sup>2</sup> AAIJ 12AL measures  $m(\Lambda_b(5912)^0) - m(\Lambda_b^0) = 292.60 \pm 0.12 \pm 0.04 \text{ MeV}$ . We have adjusted the measurement to our best value of  $m(\Lambda_b^0) = 5619.51 \pm 0.23 \text{ MeV}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

See key on page 601

Baryon Particle Listings

$\Lambda_b(5912)^0$ ,  $\Lambda_b(5920)^0$ ,  $\Sigma_b$ ,  $\Sigma_b^*$

$\Lambda_b(5912)^0$ WIDTH				
VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<0.66	90	AAIJ	12AL LHCb	$p\bar{p}$ at 7 TeV

$\Lambda_b(5912)^0$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Lambda_b^0 \pi^+ \pi^-$	seen

$\Lambda_b(5912)^0$ BRANCHING RATIOS			
$\Gamma(\Lambda_b^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE			
seen	AAIJ	12AL LHCb	$p\bar{p}$ at 7 TeV

$\Lambda_b(5912)^0$ REFERENCES			
AAIJ	12AL PRL 109 172003	R. Aaij <i>et al.</i>	(LHCb Collab.)

$\Lambda_b(5920)^0$

$J^P = \frac{3}{2}^-$

Status: \*\*\*

Quantum numbers are based on quark model expectations.

$\Lambda_b(5920)^0$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5919.81 <math>\pm</math> 0.23 OUR AVERAGE</b>			
5919.3 $\pm$ 0.5 $\pm$ 0.2	1,2 AALTONEN	13v CDF	$p\bar{p}$ at 1.96 TeV
5919.91 $\pm$ 0.09 $\pm$ 0.23	3,4 AAIJ	12AL LHCb	$p\bar{p}$ at 7 TeV
1 Measured in $\Lambda_b(5920)^0 \rightarrow \Lambda_b^0 \pi^+ \pi^-$ decays with 17.3 $^{+5.3}_{-4.6}$ events, with a significance of 3.5 sigma.			
2 AALTONEN 13v measures $m(\Lambda_b(5920)^0) - m(\Lambda_b^0) - 2m(\pi) = 20.68 \pm 0.35 \pm 0.30$ MeV. We have adjusted the measurement to our best values of $m(\Lambda_b^0) = 5619.51 \pm 0.23$ MeV and $m(\pi) = 139.57018 \pm 0.00035$ MeV. Our first error is their experiment's error and our second error is the systematic error from using our best values.			
3 Observed in $\Lambda_b(5920)^0 \rightarrow \Lambda_b^0 \pi^+ \pi^-$ decays with 52.5 $\pm$ 8.1 candidates with a significance of 10.2 sigma.			
4 AAIJ 12AL measures $m(\Lambda_b(5920)^0) - m(\Lambda_b^0) = 300.40 \pm 0.08 \pm 0.04$ MeV. We have adjusted the measurement to our best value of $m(\Lambda_b^0) = 5619.51 \pm 0.23$ MeV. Our first error is their experiment's error and our second error is the systematic error from using our best values.			

$\Lambda_b(5920)^0$ WIDTH				
VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<0.63	90	AAIJ	12AL LHCb	$p\bar{p}$ at 7 TeV

$\Lambda_b(5920)^0$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Lambda_b^0 \pi^+ \pi^-$	seen

$\Lambda_b(5920)^0$ BRANCHING RATIOS			
$\Gamma(\Lambda_b^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE			
seen	AAIJ	12AL LHCb	$p\bar{p}$ at 7 TeV

$\Lambda_b(5920)^0$ REFERENCES			
AALTONEN	13V PR D88 071101	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AAIJ	12AL PRL 109 172003	R. Aaij <i>et al.</i>	(LHCb Collab.)

$\Sigma_b$

$I(J^P) = 1(\frac{1}{2}^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

In the quark model  $\Sigma_b^+$ ,  $\Sigma_b^0$ ,  $\Sigma_b^-$  are an isotriplet ( $uub$ ,  $udb$ ,  $ddb$ ) state. The lowest  $\Sigma_b$  ought to have  $J^P = 1/2^+$ . None of  $I, J$ , or  $P$  have actually been measured.

$\Sigma_b$ MASS			
$\Sigma_b^+$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5811.3 <math>\pm</math> 0.9 <math>\pm</math> 1.7</b>	1 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5807.8 $\pm$ 2.0 $\pm$ 1.7	2 AALTONEN	07K CDF	Repl. by AALTONEN 12F

$\Sigma_b^-$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5815.5 <math>\pm</math> 0.6 <math>\pm</math> 1.7</b>	1 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5815.2 $\pm$ 1.0 $\pm$ 1.7	2 AALTONEN	07K CDF	Repl. by AALTONEN 12F

$m_{\Sigma_b^+} - m_{\Sigma_b^-}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>-4.2 <math>\pm</math> 1.1 <math>\pm</math> 0.1</b>	1 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV
1 Measured using the fully reconstructed $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ and $\Lambda_c^+ \rightarrow K^- \pi^+$ decays.			
2 Observed four $\Lambda_b^0 \pi^\pm$ resonances in the fully reconstructed decay mode $\Lambda_b^0 \rightarrow \Lambda_c^\pm \pi^-$ , where $\Lambda_c^\pm \rightarrow p K^\mp \pi^\pm$ .			

$\Sigma_b^+$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>9.7 <math>\pm</math> 3.8 <math>\pm</math> 1.2</b>	3 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV

$\Sigma_b^-$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4.9 <math>\pm</math> 3.1 <math>\pm</math> 1.1</b>	3 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV
3 Measured using the fully reconstructed $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ and $\Lambda_c^+ \rightarrow K^- \pi^+$ decays.			

$\Sigma_b$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Lambda_b^0 \pi$	dominant

$\Sigma_b$ BRANCHING RATIOS			
$\Gamma(\Lambda_b^0 \pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
VALUE			
dominant	AALTONEN	07K CDF	$p\bar{p}$ at 1.96 TeV

$\Sigma_b$ REFERENCES			
AALTONEN	12F PR D85 092011	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	07K PRL 99 202001	T. Aaltonen <i>et al.</i>	(CDF Collab.)

$\Sigma_b^*$

$I(J^P) = 1(\frac{3}{2}^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

$I, J, P$  need confirmation. Quantum numbers shown are quark-model predictions.

$\Sigma_b^*$ MASS			
$\Sigma_b^{*+}$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5832.1 <math>\pm</math> 0.7 <math>\pm</math> 1.7</b>	1 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV

$\Sigma_b^{*-}$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5835.1 <math>\pm</math> 0.6 <math>\pm</math> 1.7</b>	1 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV
$m_{\Sigma_b^{*+}} - m_{\Sigma_b^{*-}}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>-3.0 <math>\pm</math> 1.0 <math>\pm</math> 0.9</b>	1 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV

1 Measured using the fully reconstructed $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ and $\Lambda_c^+ \rightarrow K^- \pi^+$ decays.			
--	--	--	--

$\Sigma_b^*$ WIDTH			
$\Sigma_b^{*+}$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>11.5 <math>\pm</math> 2.7 <math>\pm</math> 1.0</b>	2 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV

$\Sigma_b^{*-}$ WIDTH			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>7.5 <math>\pm</math> 2.2 <math>\pm</math> 0.9</b>	2 AALTONEN	12F CDF	$p\bar{p}$ at 1.96 TeV

# Baryon Particle Listings

$\Sigma_b^*, \Xi_b^0, \Xi_b^-$

<sup>2</sup> Measured using the fully reconstructed  $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$  and  $\Lambda_c^+ \rightarrow K^- \pi^+$  decays.

$m_{\Sigma_b^*} - m_{\Sigma_b}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
$21.2^{+2.0+0.4}_{-1.9-0.3}$	<sup>3</sup> AALTONEN	07K CDF	$p\bar{p}$ at 1.96 TeV
<sup>3</sup> Observed four $\Lambda_b^0 \pi^\pm$ resonances in the fully reconstructed decay mode $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ , where $\Lambda_c^+ \rightarrow p K^- \pi^+$ . Assumes $m_{\Sigma_b^{*+}} - m_{\Sigma_b^+} = m_{\Sigma_b^{*-}} - m_{\Sigma_b^-}$			

$\Sigma_b^*$ DECAY MODES	
Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \Lambda_b^0 \pi$	dominant

$\Sigma_b^*$ BRANCHING RATIOS			
$\Gamma(\Lambda_b^0 \pi)/\Gamma_{\text{total}}$	DOCUMENT ID	TECN	COMMENT
dominant	AALTONEN	07K CDF	$p\bar{p}$ at 1.96 TeV

$\Sigma_b^*$ REFERENCES			
AALTONEN	12F	PR D85 092011	T. Aaltonen et al. (CDF Collab.)
AALTONEN	07K	PRL 99 202001	T. Aaltonen et al. (CDF Collab.)

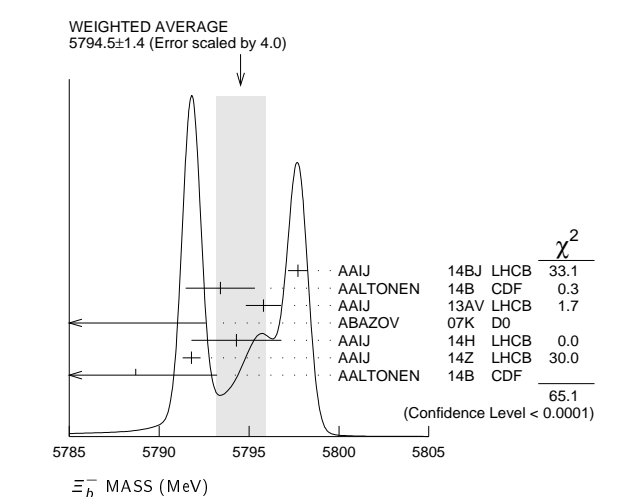
$\Xi_b^0, \Xi_b^-$

$I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$  Status: \*\*\*  
 $I, J, P$  need confirmation.

In the quark model,  $\Xi_b^0$  and  $\Xi_b^-$  are an isodoublet (*usb, dsb*) state; the lowest  $\Xi_b^0$  and  $\Xi_b^-$  ought to have  $J^P = 1/2^+$ . None of  $I, J$ , or  $P$  have actually been measured.

## $\Xi_b$ MASSES

$\Xi_b^-$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5794.5 ± 1.4 OUR AVERAGE</b>	Includes data from the datablock that follows this one. Error includes scale factor of 4.0. See the ideogram below.		
5797.72 ± 0.46 ± 0.31	<sup>1</sup> AAIJ	14BJ LHCB	$p\bar{p}$ at 7, 8 TeV
5793.4 ± 1.8 ± 0.7	<sup>2</sup> AALTONEN	14B CDF	$p\bar{p}$ at 1.96 TeV
5795.8 ± 0.9 ± 0.4	<sup>3</sup> AAIJ	13AV LHCB	$p\bar{p}$ at 7 TeV
5774 ± 11 ± 15	<sup>4</sup> ABAZOV	07K D0	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5796.7 ± 5.1 ± 1.4	<sup>5</sup> AALTONEN	11X CDF	Repl. by AALTONEN 14B
5790.9 ± 2.6 ± 0.8	<sup>6</sup> AALTONEN	09AP CDF	Repl. by AALTONEN 14B
5792.9 ± 2.5 ± 1.7	<sup>7</sup> AALTONEN	07A CDF	Repl. by AALTONEN 09AP
<sup>1</sup> Reconstructed in $\Xi_b^- \rightarrow \Xi_c^0 \pi^-, \Xi_c^0 \rightarrow p K^- K^- \pi^+$ decays. Reference $\Lambda_b^0$ mass 5619.30 ± 0.34 MeV from AAIJ 14AA.			
<sup>2</sup> Uses $\Xi_b^- \rightarrow J/\psi \Xi^-$ and $\Xi_c^0 \pi^-$ decays.			
<sup>3</sup> Measured in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays.			
<sup>4</sup> Observed in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays with 15.2 ± 4.4 <sup>+1.9</sup> <sub>-0.4</sub> candidates, a significance of 5.5 sigma.			
<sup>5</sup> Measured in $\Xi_b^- \rightarrow \Xi_c^0 \pi^-$ with 25.8 <sup>+5.5</sup> <sub>-5.2</sub> candidates.			
<sup>6</sup> Measured in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays with 66 <sup>+14</sup> <sub>-9</sub> candidates.			
<sup>7</sup> Observed in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays with 17.5 ± 4.3 candidates, a significance of 7.7 sigma.			



$\Xi_b^0$ MASS			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
The data in this block is included in the average printed for a previous datablock.			

$\Xi_b^0$ MASS			
5791.9 ± 0.5 OUR AVERAGE	Error includes scale factor of 2.1.		
5794.3 ± 2.4 ± 0.7	<sup>1</sup> AAIJ	14H LHCB	$p\bar{p}$ at 7 TeV
5791.80 ± 0.39 ± 0.31	<sup>1</sup> AAIJ	14Z LHCB	$p\bar{p}$ at 7, 8 TeV
5788.7 ± 4.3 ± 1.4	<sup>2</sup> AALTONEN	14B CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
5787.8 ± 5.0 ± 1.3	<sup>3</sup> AALTONEN	11X CDF	Repl. by AALTONEN 14B
<sup>1</sup> Uses $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ decays. The measurement comes from the mass difference of $\Xi_b^0$ and $\Lambda_b^0$ .			
<sup>2</sup> Uses $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ decays.			
<sup>3</sup> Measured in $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ with 25.3 <sup>+5.6</sup> <sub>-5.4</sub> candidates.			

$m_{\Xi_b^-} - m_{\Lambda_b^0}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>177.9 ± 0.9 OUR AVERAGE</b>	Error includes scale factor of 2.1.		
178.36 ± 0.46 ± 0.16	<sup>1</sup> AAIJ	14BJ LHCB	$p\bar{p}$ at 7, 8 TeV
176.2 ± 0.9 ± 0.1	<sup>2</sup> AAIJ	13AV LHCB	$p\bar{p}$ at 7 TeV
<sup>1</sup> Reconstructed in $\Xi_b^- \rightarrow \Xi_c^0 \pi^-, \Xi_c^0 \rightarrow p K^- K^- \pi^+$ decays. Reference $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ .			
<sup>2</sup> Reconstructed in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays.			

$m_{\Xi_b^0} - m_{\Lambda_b^0}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>172.5 ± 0.4 OUR AVERAGE</b>	Error includes scale factor of 2.1.		
174.8 ± 2.4 ± 0.5	<sup>1</sup> AAIJ	14H LHCB	$p\bar{p}$ at 7 TeV
172.44 ± 0.39 ± 0.17	<sup>1</sup> AAIJ	14Z LHCB	$p\bar{p}$ at 7, 8 TeV
<sup>1</sup> Uses $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ decays.			

$m_{\Xi_b^-} - m_{\Xi_b^0}$			
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5.9 ± 0.6 OUR AVERAGE</b>	Error includes scale factor of 2.1.		
5.92 ± 0.60 ± 0.23	<sup>1</sup> AAIJ	14BJ LHCB	$p\bar{p}$ at 7, 8 TeV
3.1 ± 5.6 ± 1.3	<sup>2</sup> AALTONEN	11X CDF	$p\bar{p}$ at 1.96 TeV
<sup>1</sup> Reconstructed in $\Xi_b^- \rightarrow \Xi_c^0 \pi^-, \Xi_c^0 \rightarrow p K^- K^- \pi^+$ decays. Uses $m(\Xi_b^0) - m(\Lambda_b^0) = 172.44 \pm 0.39 \pm 0.17$ MeV from AAIJ 14Z.			
<sup>2</sup> Derived from measurements in $\Xi_b^- \rightarrow \Xi_c^+ \pi^-$ and $\Xi_b^- \rightarrow J/\psi \Xi^-$ from AALTONEN 09AP taking correlated systematic uncertainties into account.			

**$\Xi_b$  MEAN LIFE**

"OUR EVALUATION" is an average using rescaled values of the data listed below. The average and rescaling were performed by the Heavy Flavor Averaging Group (HFAG) and are described at <http://www.slac.stanford.edu/xorg/hfag/>. The averaging/rescaling procedure takes into account correlations between the measurements and asymmetric lifetime errors.

$\Xi_b^-$ MEAN LIFE			
VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
<b>1.560 ± 0.040 OUR EVALUATION</b>	Error includes scale factor of 1.1.		
<b>1.57 ± 0.04 OUR AVERAGE</b>	Error includes scale factor of 1.1.		
1.599 ± 0.041 ± 0.022	<sup>1</sup> AAIJ	14BJ LHCB	$p\bar{p}$ at 7, 8 TeV
1.55 <sup>+0.10</sup> <sub>-0.09</sub> ± 0.03	<sup>2</sup> AAIJ	14T LHCB	$p\bar{p}$ at 7, 8 TeV
1.36 ± 0.15 ± 0.02	AALTONEN	14B CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.56 <sup>+0.27</sup> <sub>-0.25</sub> ± 0.02	<sup>3</sup> AALTONEN	09AP CDF	Repl. by AALTONEN 14B
<sup>1</sup> Reconstructed in $\Xi_b^- \rightarrow \Xi_c^0 \pi^-, \Xi_c^0 \rightarrow p K^- K^- \pi^+$ decays. Reference $\Lambda_b^0$ lifetime 1.479 ± 0.009 ± 0.010 ps from AAIJ 14U.			
<sup>2</sup> Measured in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays.			
<sup>3</sup> Measured in $\Xi_b^- \rightarrow J/\psi \Xi^-$ decays with 66 <sup>+14</sup> <sub>-9</sub> candidates.			

$\Xi_b^0$ MEAN LIFE			
VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
<b>1.464 ± 0.031 OUR EVALUATION</b>	Error includes scale factor of 1.1.		
<b>1.477 ± 0.026 ± 0.019</b>	<sup>1</sup> AAIJ	14Z LHCB	$p\bar{p}$ at 7, 8 TeV
<sup>1</sup> Uses $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ decays. The measurement comes from the value of relative lifetime of $\Xi_b^0$ to $\Lambda_b^0$ .			

$\Xi_b$ MEAN LIFE			
VALUE ( $10^{-12}$ s)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.48 <sup>+0.40</sup> <sub>-0.31</sub> ± 0.12	<sup>1</sup> ABDALLAH	05C DLPH	$e^+ e^- \rightarrow Z^0$
1.35 <sup>+0.37</sup> <sub>-0.28</sub> ± 0.15	<sup>2</sup> BUSKULIC	96T ALEP	$e^+ e^- \rightarrow Z$
1.5 <sup>+0.7</sup> <sub>-0.4</sub> ± 0.3	<sup>3</sup> ABREU	95V DLPH	Repl. by ABDALLAH 05C

See key on page 601

## Baryon Particle Listings

 $\Xi_b^0, \Xi_b^-, \Xi_b'(5935)^-, \Xi_b(5945)^0$ <sup>1</sup> Used the decay length of  $\Xi^-$  accompanied by a lepton of the same sign.<sup>2</sup> Excess  $\Xi^- \ell^-$ , impact parameters.<sup>3</sup> Excess  $\Xi^- \ell^-$ , decay lengths.

## MEAN LIFE RATIOS

 $\tau_{\Xi_b^-} / \tau_{\Lambda_b^0}$  mean life ratio

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.089 \pm 0.026 \pm 0.011</math></b>	<sup>1</sup> AAIJ	14BJ LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Reconstructed in  $\Xi_b^- \rightarrow \Xi_c^0 \pi^-, \Xi_c^0 \rightarrow p K^- K^- \pi^+$  decays. Reference  $\Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ . $\tau_{\Xi_b^-} / \tau_{\Xi_b^0}$  mean life ratio

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>1.083 \pm 0.032 \pm 0.016</math></b>	<sup>1</sup> AAIJ	14BJ LHCb	$pp$ at 7, 8 TeV

<sup>1</sup> Reconstructed in  $\Xi_b^- \rightarrow \Xi_c^0 \pi^-, \Xi_c^0 \rightarrow p K^- K^- \pi^+$  decays. Uses  $\Xi_b^0$  measurements from AAIJ 14Z. $\Xi_b$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor / Confidence level
$\Gamma_1 \quad \Xi_b^- \rightarrow \Xi^- \ell^- \bar{\nu}_\ell X \times B(\bar{b} \rightarrow \Xi_b)$	$(3.9 \pm 1.2) \times 10^{-4}$	S=1.4
$\Gamma_2 \quad \Xi_b^- \rightarrow J/\psi \Xi^- \times B(b \rightarrow \Xi_b^-)$	$(1.02^{+0.26}_{-0.21}) \times 10^{-5}$	
$\Gamma_3 \quad \Xi_b^0 \rightarrow p D^0 K^- \times B(\bar{b} \rightarrow \Xi_b)$	$(1.7 \pm 0.6) \times 10^{-6}$	
$\Gamma_4 \quad \Xi_b^0 \rightarrow p \bar{K}^0 \pi^- \times B(\bar{b} \rightarrow \Xi_b)/B(\bar{b} \rightarrow B^0)$	$< 1.6 \times 10^{-6}$	CL=90%
$\Gamma_5 \quad \Xi_b^0 \rightarrow p K^0 K^- \times B(\bar{b} \rightarrow \Xi_b)/B(\bar{b} \rightarrow B^0)$	$< 1.1 \times 10^{-6}$	CL=90%
$\Gamma_6 \quad \Xi_b^0 \rightarrow \Lambda_c^+ K^- \times B(\bar{b} \rightarrow \Xi_b)$	$(6 \pm 4) \times 10^{-7}$	
$\Gamma_7 \quad \Xi_b^- \rightarrow \Lambda_b^0 \pi^- \times B(b \rightarrow \Xi_b^-)/B(b \rightarrow \Lambda_b^0)$	$(5.7 \pm 2.0) \times 10^{-4}$	

 $\Xi_b$  BRANCHING RATIOS

$\Gamma(\Xi^- \ell^- \bar{\nu}_\ell X \times B(\bar{b} \rightarrow \Xi_b))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>3.9 \pm 1.2</math> OUR AVERAGE</b>	Error includes scale factor of 1.4.
$3.0 \pm 1.0 \pm 0.3$	ABDALLAH 05c DLPH $e^+ e^- \rightarrow Z^0$
$5.4 \pm 1.1 \pm 0.8$	BUSKULIC 96T ALEP Excess $\Xi^- \ell^-$ over $\Xi^- \ell^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •	
$5.9 \pm 2.1 \pm 1.0$	ABREU 95V DLPH Repl. by ABDALLAH 05c

$\Gamma(J/\psi \Xi^- \times B(b \rightarrow \Xi_b^-))/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>0.102^{+0.026}_{-0.021}</math> OUR AVERAGE</b>	
$0.098^{+0.023}_{-0.016} \pm 0.014$	<sup>1</sup> AALTONEN 09AP CDF $p\bar{p}$ at 1.96 TeV
$0.16 \pm 0.07 \pm 0.02$	<sup>2</sup> ABAZOV 07K D0 $p\bar{p}$ at 1.96 TeV

<sup>1</sup> AALTONEN 09AP reports  $[\Gamma(\Xi_b^- \rightarrow J/\psi \Xi^- \times B(b \rightarrow \Xi_b^-))/\Gamma_{\text{total}}] / [B(\Lambda_b^0 \rightarrow J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0))] = 0.167^{+0.037}_{-0.025} \pm 0.012$  which we multiply by our best value  $B(\Lambda_b^0 \rightarrow J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0)) = (5.8 \pm 0.8) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.<sup>2</sup> ABAZOV 07K reports  $[\Gamma(\Xi_b^- \rightarrow J/\psi \Xi^- \times B(b \rightarrow \Xi_b^-))/\Gamma_{\text{total}}] / [B(\Lambda_b^0 \rightarrow J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0))] = 0.28 \pm 0.09^{+0.09}_{-0.08}$  which we multiply by our best value  $B(\Lambda_b^0 \rightarrow J/\psi(1S) \Lambda \times B(b \rightarrow \Lambda_b^0)) = (5.8 \pm 0.8) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p D^0 K^- \times B(\bar{b} \rightarrow \Xi_b))/\Gamma_{\text{total}}$	$\Gamma_3/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>(1.8 \pm 0.4 \pm 0.4) \times 10^{-6}</math></b>	<sup>1</sup> AAIJ 14H LHCb $pp$ at 7 TeV

<sup>1</sup> AAIJ 14H reports  $[\Gamma(\Xi_b^0 \rightarrow p D^0 K^- \times B(\bar{b} \rightarrow \Xi_b))/\Gamma_{\text{total}}] / [B(\bar{b} \rightarrow b\text{-baryon})] / [B(\Lambda_b^0 \rightarrow p D^0 K^-)] = 0.44 \pm 0.09 \pm 0.06$  which we multiply by our best values  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ ,  $B(\Lambda_b^0 \rightarrow p D^0 K^-) = (4.7 \pm 0.8) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

$\Gamma(p \bar{K}^0 \pi^- \times B(\bar{b} \rightarrow \Xi_b)/B(\bar{b} \rightarrow B^0))/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT
<b><math>&lt; 1.6 \times 10^{-6}</math></b>	90 AAIJ 14Q LHCb $pp$ at 7 TeV

$\Gamma(p K^0 K^- \times B(\bar{b} \rightarrow \Xi_b)/B(\bar{b} \rightarrow B^0))/\Gamma_{\text{total}}$	$\Gamma_5/\Gamma$
VALUE CL%	DOCUMENT ID TECN COMMENT
<b><math>&lt; 1.1 \times 10^{-6}</math></b>	90 AAIJ 14Q LHCb $pp$ at 7 TeV

$\Gamma(\Lambda_c^+ K^- \times B(\bar{b} \rightarrow \Xi_b))/\Gamma(p D^0 K^- \times B(\bar{b} \rightarrow \Xi_b))$	$\Gamma_6/\Gamma_3$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.35 \pm 0.19 \pm 0.02</math></b>	<sup>1</sup> AAIJ 14H LHCb $pp$ at 7 TeV

<sup>1</sup> AAIJ 14H reports  $[\Gamma(\Xi_b^0 \rightarrow \Lambda_c^+ K^- \times B(\bar{b} \rightarrow \Xi_b))/\Gamma(\Xi_b^0 \rightarrow p D^0 K^- \times B(\bar{b} \rightarrow \Xi_b))] \times [B(\Lambda_c^+ \rightarrow p K^- \pi^+)] / [B(D^0 \rightarrow K^- \pi^+)] = 0.57 \pm 0.22 \pm 0.21$  which we multiply or divide by our best values  $B(\Lambda_c^+ \rightarrow p K^- \pi^+) = (6.35 \pm 0.33) \times 10^{-2}$ ,  $B(D^0 \rightarrow K^- \pi^+) = (3.93 \pm 0.04) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best values.

$\Gamma(\Lambda_b^0 \pi^- \times B(b \rightarrow \Xi_b^-)/B(b \rightarrow \Lambda_b^0))/\Gamma_{\text{total}}$	$\Gamma_7/\Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID TECN COMMENT
<b><math>5.7 \pm 1.8^{+0.8}_{-0.9}</math></b>	<sup>1</sup> AAIJ 15BA LHCb $pp$ at 7, 8 TeV

<sup>1</sup> A signal is reported with a significance of 3.2 standard deviations in the decay chain of  $\Xi_b^- \rightarrow \Lambda_b^0 \pi^-, \Lambda_b^0 \rightarrow \Lambda_c^+ \pi^-$ , and  $\Lambda_c^+ \rightarrow p K^- \pi^+$ . $\Xi_b$  REFERENCES

AAIJ 15BA PRL 115 241801	R. Aaij et al.	(LHCb Collab.)
AAIJ 14AA PRL 112 202001	R. Aaij et al.	(LHCb Collab.)
AAIJ 14BJ PRL 113 242002	R. Aaij et al.	(LHCb Collab.)
AAIJ 14H PR D89 032001	R. Aaij et al.	(LHCb Collab.)
AAIJ 14Q JHEP 1404 087	R. Aaij et al.	(LHCb Collab.)
AAIJ 14T PL B736 154	R. Aaij et al.	(LHCb Collab.)
AAIJ 14U PL B734 122	R. Aaij et al.	(LHCb Collab.)
AAIJ 14Z PRL 113 032001	R. Aaij et al.	(LHCb Collab.)
AALTONEN 14B PR D89 072014	T. Aaltonen et al.	(CDF Collab.)
AAIJ 13AV PRL 110 182001	R. Aaij et al.	(LHCb Collab.)
AALTONEN 11X PRL 107 102001	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 09AP PR D80 072003	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 07A PRL 99 052002	T. Aaltonen et al.	(CDF Collab.)
ABAZOV 07K PRL 99 052001	V.M. Abazov et al.	(D0 Collab.)
ABDALLAH 05C EPJ C44 239	J. Abdallah et al.	(DELPHI Collab.)
BUSKULIC 96T PL B384 449	D. Buskulic et al.	(ALEPH Collab.)
ABREU 95V ZPHY C68 541	P. Abreu et al.	(DELPHI Collab.)

 $\Xi_b'(5935)^-$ 

$$J^P = \frac{1}{2}^+$$

Status: \*\*\*

 $\Xi_b'(5935)^-$  MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>5935.02 \pm 0.02 \pm 0.05</math></b>	<sup>1</sup> AAIJ 15H LHCb		$pp$ at 7, 8 TeV

<sup>1</sup> Not independent of the mass difference measurement below. Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$  and  $\Xi_c^+ \rightarrow p K^- \pi^+$ .

<sup>1</sup> Not independent of the mass difference measurement below. Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$  and  $\Xi_c^+ \rightarrow p K^- \pi^+$ .

$$m_{\Xi_b'(5935)^-} - m_{\Xi_b^0} - m_{\pi^-}$$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b><math>3.653 \pm 0.018 \pm 0.006</math></b>	<sup>2</sup> AAIJ 15H LHCb		$pp$ at 7, 8 TeV

<sup>2</sup>Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$  and  $\Xi_c^+ \rightarrow p K^- \pi^+$ .

<sup>2</sup> Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$  and  $\Xi_c^+ \rightarrow p K^- \pi^+$ . $\Xi_b'(5935)^-$  WIDTH

VALUE (MeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.08</b>	95	<sup>3</sup> AAIJ 15H LHCb		$pp$ at 7, 8 TeV

<sup>3</sup> Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi^+ \pi^-$  and  $\Xi_b^+ \rightarrow p K^- \pi^+$ .

<sup>3</sup> Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$  and  $\Xi_c^+ \rightarrow p K^- \pi^+$ . $\Xi_b'(5935)^-$  DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Xi_b^0 \pi^- \times B(\bar{b} \rightarrow \Xi_b'(5935)^-)/B(\bar{b} \rightarrow \Xi_b^0)$	$(11.8 \pm 1.8) \%$

 $\Xi_b'(5935)^-$  BRANCHING RATIOS

$\Gamma(\Xi_b^0 \pi^- \times B(\bar{b} \rightarrow \Xi_b'(5935)^-)/B(\bar{b} \rightarrow \Xi_b^0))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID TECN COMMENT
<b><math>0.118 \pm 0.017 \pm 0.007</math></b>	<sup>4</sup> AAIJ 15H LHCb $pp$ at 7, 8 TeV

<sup>4</sup> Observed in  $\Xi_b^0 \pi^-$  channel with  $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$  and  $\Xi_c^+ \rightarrow p K^- \pi^+$ . $\Xi_b'(5935)^-$  REFERENCES

AAIJ 15H PRL 114 062004	R. Aaij et al.	(LHCb Collab.)
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 $\Xi_b(5945)^0$ 

$$J^P = \frac{3}{2}^+$$

Status: \*\*\*

Quantum numbers are based on quark model expectations.

# Baryon Particle Listings

## $\Xi_b(5945)^0$ , $\Xi_b^{*}(5955)^-$ , $\Omega_b^-$

### $\Xi_b(5945)^0$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5948.9±0.8±1.4</b>	<sup>1</sup> CHATRCHYAN12s	CMS	$pp$ at 7 TeV, 5.3 fb <sup>-1</sup>
<sup>1</sup> CHATRCHYAN 12s measures $m(\Xi_b(5945)^0) - m(\Xi_b^-) - m(\pi^+) = 14.84 \pm 0.74 \pm 0.28$ MeV. We have adjusted the measurement to our best values of $m(\Xi_b^-) = 5794.5 \pm 1.4$ MeV, $m(\pi^+) = 139.57018 \pm 0.00035$ MeV. Our first error is their experiment's error and our second error is the systematic error from using our best values.			

### $\Xi_b(5945)^0$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>2.1±1.7</b>	<sup>2</sup> CHATRCHYAN12s	CMS	$pp$ at 7 TeV, 5.3 fb <sup>-1</sup>
<sup>2</sup> Systematic uncertainty not evaluated.			

### $\Xi_b(5945)^0$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Xi_b^- \pi^+$	seen

### $\Xi_b(5945)^0$ BRANCHING RATIOS

$\Gamma(\Xi_b^- \pi^+)/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
seen	CHATRCHYAN12s	CMS	$pp$ at 7 TeV, 5.3 fb <sup>-1</sup>

### $\Xi_b(5945)^0$ REFERENCES

CHATRCHYAN 12s	PRL 108 252002	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
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$\Xi_b^{*}(5955)^-$	$J^P = \frac{3}{2}^+$	Status: ***
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### $\Xi_b^{*}(5955)^-$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>5955.33±0.12±0.05</b>	<sup>1</sup> AAIJ	15H LHCb	$pp$ at 7, 8 TeV
<sup>1</sup> Not independent of the mass difference measurement below. Observed in $\Xi_b^0 \pi^-$ channel with $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ .			

### $m_{\Xi_b^{*}(5955)^-} - m_{\Xi_b^0} - m_{\pi^-}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>23.96±0.12±0.06</b>	<sup>2</sup> AAIJ	15H LHCb	$pp$ at 7, 8 TeV
<sup>2</sup> Observed in $\Xi_b^0 \pi^-$ channel with $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ .			

### $\Xi_b^{*}(5955)^-$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>1.65±0.31±0.10</b>	<sup>3</sup> AAIJ	15H LHCb	$pp$ at 7, 8 TeV
<sup>3</sup> Observed in $\Xi_b^0 \pi^-$ channel with $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ .			

### $\Xi_b^{*}(5955)^-$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad \Xi_b^0 \pi^- \times \text{B}(\bar{b} \rightarrow \Xi_b^{*}(5955)^-)/\text{B}(\bar{b} \rightarrow \Xi_b^0)$	(20.7±3.5) %

### $\Xi_b^{*}(5955)^-$ BRANCHING RATIOS

$\Gamma(\Xi_b^0 \pi^- \times \text{B}(\bar{b} \rightarrow \Xi_b^{*}(5955)^-)/\text{B}(\bar{b} \rightarrow \Xi_b^0))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.207±0.032±0.015</b>	<sup>4</sup> AAIJ	15H LHCb	$pp$ at 7, 8 TeV
<sup>4</sup> Observed in $\Xi_b^0 \pi^-$ channel with $\Xi_b^0 \rightarrow \Xi_c^+ \pi^-$ and $\Xi_c^+ \rightarrow p K^- \pi^+$ .			

### $\Xi_b^{*}(5955)^-$ REFERENCES

AAIJ	15H	PRL 114 062004	R. Aaij <i>et al.</i>	(LHCb Collab.)
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$\Omega_b^-$	$I(J^P) = 0(\frac{1}{2}^+)$ Status: *** $I, J, P$ need $p$ confirmation.
In the quark model $\Omega_b^-$ is $ssb$ ground state. None of its quantum numbers has been measured.	

### $\Omega_b^-$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>6046.4±1.9 OUR AVERAGE</b>			
6047.5±3.8±0.6	<sup>1</sup> AALTONEN	14B CDF	$p\bar{p}$ at 1.96 TeV
6046.0±2.2±0.5	<sup>2</sup> AAIJ	13AV LHCb	$pp$ at 7 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
6054.4±6.8±0.9	<sup>3</sup> AALTONEN	09AP CDF	Repl. by AALTONEN 14b
6165±10±13	<sup>4</sup> ABAZOV	08AL D0	$p\bar{p}$ at 1.96 TeV

- <sup>1</sup> Uses  $\Omega_b^- \rightarrow J/\psi \Omega^-$  and  $\Omega_c^0 \pi^-$  decays, with the first evidence for  $\Omega_b^- \rightarrow \Omega_c^0 \pi^-$  at 3.3  $\sigma$  significance.
- <sup>2</sup> Measured in  $\Omega_b^- \rightarrow J/\psi \Omega^-$  with  $19 \pm 5$  events.
- <sup>3</sup> Observed in  $\Omega_b^- \rightarrow J/\psi \Omega^-$  decays with  $16^{+6}_{-4}$  candidates, a significance of 5.5 sigma from a combined mass-lifetime fit.
- <sup>4</sup> Observed in  $\Omega_b^- \rightarrow J/\psi \Omega^-$  decays with  $17.8 \pm 4.9 \pm 0.8$  candidates, a significance of 5.4 sigma.

### $m_{\Omega_b^-} - m_{\Lambda_b^0}$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>426.4±2.2±0.4</b>	AAIJ	13AV LHCb	$pp$ at 7 TeV

### $\Omega_b^-$ MEAN LIFE

VALUE (10 <sup>-12</sup> s)	DOCUMENT ID	TECN	COMMENT
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### 1.57<sup>+0.23</sup><sub>-0.20</sub> OUR EVALUATION

### 1.57<sup>+0.24</sup><sub>-0.19</sub> OUR AVERAGE

1.54 <sup>+0.26</sup> <sub>-0.21</sub> ±0.05	<sup>1</sup> AAIJ	14T LHCb	$pp$ at 7, 8 TeV
1.66 <sup>+0.53</sup> <sub>-0.40</sub> ±0.02	<sup>1</sup> AALTONEN	14B CDF	$p\bar{p}$ at 1.96 TeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.13 <sup>+0.53</sup> <sub>-0.40</sub> ±0.02	<sup>2</sup> AALTONEN	09AP CDF	Repl. by AALTONEN 14b

- <sup>1</sup> Measured in  $\Omega_b^- \rightarrow J/\psi \Omega^-$  decays.
- <sup>2</sup> Observed in  $\Omega_b^- \rightarrow J/\psi \Omega^-$  decays with  $16^{+6}_{-4}$  candidates, a significance of 5.5 sigma from a combined mass-lifetime fit.

### $\Omega_b^-$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \quad J/\psi \Omega^- \times \text{B}(b \rightarrow \Omega_b)$	(2.9 <sup>+1.1</sup> <sub>-0.8</sub> ) × 10 <sup>-6</sup>

### $\Omega_b^-$ BRANCHING RATIOS

$\Gamma(J/\psi \Omega^- \times B(b \rightarrow \Omega_b))/\Gamma_{\text{total}}$	$\Gamma_1/\Gamma$		
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
$0.029^{+0.011}_{-0.008}$ OUR AVERAGE			

0.026 <sup>+0.010</sup> <sub>-0.007</sub> ±0.004	<sup>1</sup> AALTONEN	09AP CDF	$p\bar{p}$ at 1.96 TeV
0.08±0.04±0.02	<sup>2</sup> ABAZOV	08AL D0	$p\bar{p}$ at 1.96 TeV

- <sup>1</sup> AALTONEN 09AP reports  $[\Gamma(\Omega_b^- \rightarrow J/\psi \Omega^- \times \text{B}(b \rightarrow \Omega_b))/\Gamma_{\text{total}}] / [\text{B}(\Lambda_b^0 \rightarrow J/\psi(1S) \Lambda \times \text{B}(b \rightarrow \Lambda_b^0))] = 0.045^{+0.017}_{-0.012} \pm 0.004$  which we multiply by our best value  $\text{B}(\Lambda_b^0 \rightarrow J/\psi(1S) \Lambda \times \text{B}(b \rightarrow \Lambda_b^0)) = (5.8 \pm 0.8) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.
- <sup>2</sup> ABAZOV 08AL reports  $[\Gamma(\Omega_b^- \rightarrow J/\psi \Omega^- \times \text{B}(b \rightarrow \Omega_b))/\Gamma_{\text{total}}] / [\text{B}(\Xi_b^- \rightarrow J/\psi \Xi^- \times \text{B}(b \rightarrow \Xi_b^-))] = 0.80 \pm 0.32^{+0.14}_{-0.22}$  which we multiply by our best value  $\text{B}(\Xi_b^- \rightarrow J/\psi \Xi^- \times \text{B}(b \rightarrow \Xi_b^-)) = (1.02^{+0.26}_{-0.21}) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

### $\Omega_b^-$ REFERENCES

AAIJ	14T	PL B736 154	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	14B	PR D89 072014	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AAIJ	13AV	PRL 110 182001	R. Aaij <i>et al.</i>	(LHCb Collab.)
AALTONEN	09AP	PR D80 072003	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	08AL	PRL 101 232002	V.M. Abazov <i>et al.</i>	(D0 Collab.)



See key on page 601

# Baryon Particle Listings

## $b$ -baryon ADMIXTURE ( $\Lambda_b$ , $\Xi_b$ , $\Sigma_b$ , $\Omega_b$ )

### $b$ -baryon ADMIXTURE ( $\Lambda_b$ , $\Xi_b$ , $\Sigma_b$ , $\Omega_b$ )

#### $b$ -baryon ADMIXTURE MEAN LIFE

Each measurement of the  $b$ -baryon mean life is an average over an admixture of various  $b$  baryons which decay weakly. Different techniques emphasize different admixtures of produced particles, which could result in a different  $b$ -baryon mean life. More  $b$ -baryon flavor specific channels are not included in the measurement.

VALUE (10 <sup>-12</sup> s)	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.218 <sup>+0.130</sup> <sub>-0.115</sub> ± 0.042		<sup>1</sup> ABAZOV	07s D0	Repl. by ABAZOV 12u
1.22 <sup>+0.22</sup> <sub>-0.18</sub> ± 0.04		<sup>1</sup> ABAZOV	05c D0	Repl. by ABAZOV 07s
1.16 ± 0.20 ± 0.08		<sup>2</sup> ABREU	99w DLPH	$e^+e^- \rightarrow Z$
1.19 ± 0.14 ± 0.07		<sup>3</sup> ABREU	99w DLPH	$e^+e^- \rightarrow Z$
1.14 ± 0.08 ± 0.04		<sup>4</sup> ABREU	99w DLPH	$e^+e^- \rightarrow Z$
1.11 <sup>+0.19</sup> <sub>-0.18</sub> ± 0.05		<sup>5</sup> ABREU	99w DLPH	$e^+e^- \rightarrow Z$
1.29 <sup>+0.24</sup> <sub>-0.22</sub> ± 0.06		<sup>5</sup> ACKERSTAFF	98g OPAL	$e^+e^- \rightarrow Z$
1.20 ± 0.08 ± 0.06		<sup>6</sup> BARATE	98d ALEP	$e^+e^- \rightarrow Z$
1.21 ± 0.11		<sup>5</sup> BARATE	98d ALEP	$e^+e^- \rightarrow Z$
1.32 ± 0.15 ± 0.07		<sup>7</sup> ABE	96m CDF	$p\bar{p}$ at 1.8 TeV
1.46 <sup>+0.22</sup> <sub>-0.21</sub> ± 0.09		ABREU	96d DLPH	Repl. by ABREU 99w
1.10 <sup>+0.19</sup> <sub>-0.17</sub> ± 0.09		<sup>5</sup> ABREU	96d DLPH	$e^+e^- \rightarrow Z$
1.16 ± 0.11 ± 0.06		<sup>5</sup> AKERS	96 OPAL	$e^+e^- \rightarrow Z$
1.27 <sup>+0.35</sup> <sub>-0.29</sub> ± 0.09		ABREU	95s DLPH	Repl. by ABREU 99w
1.05 <sup>+0.12</sup> <sub>-0.11</sub> ± 0.09	290	BUSKULIC	95l ALEP	Repl. by BARATE 98d
1.04 <sup>+0.48</sup> <sub>-0.38</sub> ± 0.10	11	<sup>8</sup> ABREU	93f DLPH	Excess $\Lambda\ell^-$ , decay lengths
1.05 <sup>+0.23</sup> <sub>-0.20</sub> ± 0.08	157	<sup>9</sup> AKERS	93 OPAL	Excess $\Lambda\ell^-$ , decay lengths
1.12 <sup>+0.32</sup> <sub>-0.29</sub> ± 0.16	101	<sup>10</sup> BUSKULIC	92i ALEP	Excess $\Lambda\ell^-$ , impact parameters

<sup>1</sup> Measured mean life using fully reconstructed  $\Lambda_b^0 \rightarrow J/\psi\Lambda$  decays.

<sup>2</sup> Measured using  $\Lambda\ell^-$  decay length.

<sup>3</sup> Measured using  $p\ell^-$  decay length.

<sup>4</sup> This ABREU 99w result is the combined result of the  $\Lambda\ell^-$ ,  $p\ell^-$ , and excess  $\Lambda\mu^-$  impact parameter measurements.

<sup>5</sup> Measured using  $\Lambda_C\ell^-$  and  $\Lambda\ell^+\ell^-$ .

<sup>6</sup> Measured using the excess of  $\Lambda\ell^-$ , lepton impact parameter.

<sup>7</sup> Measured using  $\Lambda_C\ell^-$ .

<sup>8</sup> ABREU 93f superseded by ABREU 96d.

<sup>9</sup> AKERS 93 superseded by AKERS 96.

<sup>10</sup> BUSKULIC 92i superseded by BUSKULIC 95L.

#### $b$ -baryon ADMIXTURE DECAY MODES ( $\Lambda_b$ , $\Xi_b$ , $\Sigma_b$ , $\Omega_b$ )

These branching fractions are actually an average over weakly decaying  $b$ -baryons weighted by their production rates at the LHC, LEP, and Tevatron, branching ratios, and detection efficiencies. They scale with the  $b$ -baryon production fraction  $B(b \rightarrow b\text{-baryon})$ .

The branching fractions  $B(b\text{-baryon} \rightarrow \Lambda\ell^-\bar{\nu}_\ell\text{anything})$  and  $B(\Lambda_b^0 \rightarrow \Lambda_C^+\ell^-\bar{\nu}_\ell\text{anything})$  are not pure measurements because the underlying measured products of these with  $B(b \rightarrow b\text{-baryon})$  were used to determine  $B(b \rightarrow b\text{-baryon})$ , as described in the note "Production and Decay of  $b$ -Flavored Hadrons."

For inclusive branching fractions, e.g.,  $B \rightarrow D^\pm\text{anything}$ , the values usually are multiplicities, not branching fractions. They can be greater than one.

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $p\mu^-\bar{\nu}$ anything	( 5.5 <sup>+2.2</sup> <sub>-1.9</sub> ) %
$\Gamma_2$ $p\ell\bar{\nu}_\ell$ anything	( 5.3 ± 1.2 ) %
$\Gamma_3$ $p$ anything	( 66 ± 21 ) %
$\Gamma_4$ $\Lambda\ell^-\bar{\nu}_\ell$ anything	( 3.6 ± 0.6 ) %
$\Gamma_5$ $\Lambda\ell^+\nu_\ell$ anything	( 3.0 ± 0.8 ) %
$\Gamma_6$ $\Lambda$ anything	( 37 ± 7 ) %
$\Gamma_7$ $\Xi^-\ell^-\bar{\nu}_\ell$ anything	( 6.2 ± 1.6 ) × 10 <sup>-3</sup>

#### $b$ -baryon ADMIXTURE ( $\Lambda_b$ , $\Xi_b$ , $\Sigma_b$ , $\Omega_b$ ) BRANCHING RATIOS

$\Gamma(p\mu^- \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
$0.055^{+0.021}_{-0.017} \pm 0.008$	125	<sup>11</sup> ABREU	95s DLPH	$e^+e^- \rightarrow Z$	

<sup>11</sup> ABREU 95s reports  $[\Gamma(b\text{-baryon} \rightarrow p\mu^-\bar{\nu}\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = 0.0049 \pm 0.0011^{+0.0015}_{-0.0011}$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p\ell\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}$	$\Gamma_2/\Gamma$		
VALUE	DOCUMENT ID	TECN	COMMENT
$0.053\pm0.009\pm0.008$	<sup>12</sup> BARATE	98v ALEP	$e^+e^- \rightarrow Z$

<sup>12</sup> BARATE 98v reports  $[\Gamma(b\text{-baryon} \rightarrow p\ell\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = (4.72 \pm 0.66 \pm 0.44) \times 10^{-3}$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(p\ell\bar{\nu}_\ell\text{anything})/\Gamma(p\text{anything})$				$\Gamma_2/\Gamma_3$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.080\pm0.012\pm0.014$	BARATE	98v ALEP	$e^+e^-\rightarrow Z$	

$\Gamma(\Lambda\ell^-\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}$	$\Gamma_4/\Gamma$
<p>The values and averages in this section serve only to show what values result if one assumes our <math>B(b \rightarrow b\text{-baryon})</math>. They cannot be thought of as measurements since the underlying product branching fractions were also used to determine <math>B(b \rightarrow b\text{-baryon})</math> as described in the note on "Production and Decay of <math>b</math>-Flavored Hadrons."</p>	

• • • We do not use the following data for averages, fits, limits, etc. • • •

seen 157 <sup>17</sup> AKERS 93 OPAL Excess of  $\Lambda\ell^-$  over  $\Lambda\ell^+$

0.079 ± 0.023 ± 0.011 101 <sup>18</sup> BUSKULIC 92i ALEP Excess of  $\Lambda\ell^-$  over  $\Lambda\ell^+$

<sup>13</sup> BARATE 98d reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\ell^-\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = 0.00326 \pm 0.00016 \pm 0.00039$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Measured using the excess of  $\Lambda\ell^-$ , lepton impact parameter.

<sup>14</sup> AKERS 96 reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\ell^-\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = 0.00291 \pm 0.00023 \pm 0.00025$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>15</sup> ABREU 95s reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\ell^-\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = 0.0030 \pm 0.0006 \pm 0.0004$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>16</sup> BUSKULIC 95L reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\ell^-\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = 0.0061 \pm 0.0006 \pm 0.0010$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>17</sup> AKERS 93 superseded by AKERS 96.

<sup>18</sup> BUSKULIC 92i reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\ell^-\bar{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [B(\bar{b} \rightarrow b\text{-baryon})] = 0.0070 \pm 0.0010 \pm 0.0018$  which we divide by our best value  $B(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Superseded by BUSKULIC 95L.

$\Gamma(\Lambda\ell^+\nu_\ell\text{anything})/\Gamma(\Lambda\text{anything})$				$\Gamma_5/\Gamma_6$
VALUE	DOCUMENT ID	TECN	COMMENT	
$0.080\pm0.012\pm0.008$	ABBIENDI	99L OPAL	$e^+e^-\rightarrow Z$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.070 ± 0.012 ± 0.007 ACKERSTAFF 97N OPAL Repl. by ABBI-  
ENDI 99L

$\Gamma(\Lambda\text{anything})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
0.37±0.07 OUR AVERAGE				

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.44 ± 0.07 ± 0.06 <sup>21</sup> ACKERSTAFF 97N OPAL Repl. by ABBI-  
ENDI 99L

# Baryon Particle Listings

## *b*-baryon ADMIXTURE ( $\Lambda_b, \Xi_b, \Sigma_b, \Omega_b$ )

<sup>19</sup>ABBIENDI 99L reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.035 \pm 0.0032 \pm 0.0035$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>20</sup>ABREU 95c reports  $0.28^{+0.17}_{-0.12}$  from a measurement of  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})]$  assuming  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = 0.08 \pm 0.02$ , which we rescale to our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>21</sup>ACKERSTAFF 97N reports  $[\Gamma(b\text{-baryon} \rightarrow \Lambda\text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.0393 \pm 0.0046 \pm 0.0037$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

$\Gamma(\Xi^-\ell^-\overline{\nu}_\ell\text{anything})/\Gamma_{\text{total}}$					$\Gamma_7/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.0062±0.0016 OUR AVERAGE</b>					
0.0061±0.0015±0.0009	<sup>22</sup> BUSKULIC	96T	ALEP	Excess $\Xi^-\ell^+$ over $\Xi^-\ell^+$	
0.0066±0.0026±0.0010	<sup>23</sup> ABREU	95v	DLPH	Excess $\Xi^-\ell^-$ over $\Xi^-\ell^+$	

<sup>22</sup>BUSKULIC 96T reports  $[\Gamma(b\text{-baryon} \rightarrow \Xi^-\ell^-\overline{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.00054 \pm 0.00011 \pm 0.00008$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

<sup>23</sup>ABREU 95v reports  $[\Gamma(b\text{-baryon} \rightarrow \Xi^-\ell^-\overline{\nu}_\ell\text{anything})/\Gamma_{\text{total}}] \times [\text{B}(\bar{b} \rightarrow b\text{-baryon})] = 0.00059 \pm 0.00021 \pm 0.0001$  which we divide by our best value  $\text{B}(\bar{b} \rightarrow b\text{-baryon}) = (8.9 \pm 1.3) \times 10^{-2}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value.

## *b*-baryon ADMIXTURE ( $\Lambda_b, \Xi_b, \Sigma_b, \Omega_b$ ) REFERENCES

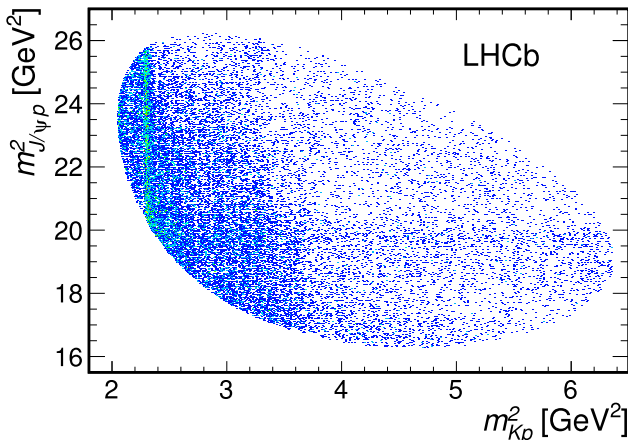
ABAZOV	12U	PR D85 112003	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	07S	PRL 99 142001	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	05C	PRL 94 102001	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABBIENDI	99L	EPJ C9 1	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	99W	EPJ C10 185	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	98G	PL B426 161	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BARATE	98D	EPJ C2 197	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	98V	EPJ C5 205	R. Barate <i>et al.</i>	(ALEPH Collab.)
ACKERSTAFF	97N	ZPHY C74 423	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ABE	96M	PRL 77 1439	F. Abe <i>et al.</i>	(CDF Collab.)
ABREU	96D	ZPHY C71 199	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	96	ZPHY C69 195	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	96T	PL B384 449	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ABREU	95C	PL B347 447	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95S	ZPHY C68 375	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95V	ZPHY C68 541	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BUSKULIC	95L	PL B357 685	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ABREU	93F	PL B311 379	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	93	PL B316 435	R. Akers <i>et al.</i>	(OPAL Collab.)
BUSKULIC	92I	PL B297 449	D. Buskulic <i>et al.</i>	(ALEPH Collab.)

## EXOTIC BARYONS

### PENTAQUARKS

Written March 2016 by M. Karliner (Tel Aviv U.), T. Skwarnicki (Syracuse U.)

Experimental searches for pentaquark hadrons comprised of light flavors have a long and vivid history. No undisputed candidates have been found in 50 years. The first wave of observations of pentaquark candidates containing a strange antiquark occurred in the early seventies, see e.g. a review in the 1976 edition of Particle Data Group listings for  $Z_0(1780)$ ,  $Z_0(1865)$  and  $Z_1(1900)$  [1]. The last mention of these candidates can be found in the 1992 edition [2] with the perhaps prophetic comment “the results permit no definite conclusion - the same story for 20 years. [...] The skepticism about baryons not made of three quarks, and lack of any experimental activity in this area, make it likely that another 20 years will pass before the issue is decided.” A decade later, a second wave of observations occurred, possibly motivated by specific theoretical predictions for their existence [3–5]. The evidence for pentaquarks was based on observations of peaks in the invariant mass distributions of their decay products. More data, or more sensitive experiments did not confirm these claims [6]. In the last mention of the best known candidate from that period,  $\Theta(1540)^+$ , the 2006 Particle Data Group listing [7] included a statement: “The conclusion that pentaquarks in general, and that  $\Theta^+$ , in particular, do not exist, appears compelling.” which well reflected the prevailing mood in the particle physics community until a study of  $\Lambda_b^0 \rightarrow J\psi p K^-$  ( $J\psi \rightarrow \mu^+\mu^-$ ) decays by LHCb [8] (charge conjugate modes are implied). In addition to many excitations of the  $\Lambda$  baryon (hereafter denoted as  $\Lambda^*$  resonances) decaying to  $K^-p$ , these data contain a narrow peak in the  $J\psi p$  mass distribution, which is evident as a horizontal band in the Dalitz plot (Fig. 1).



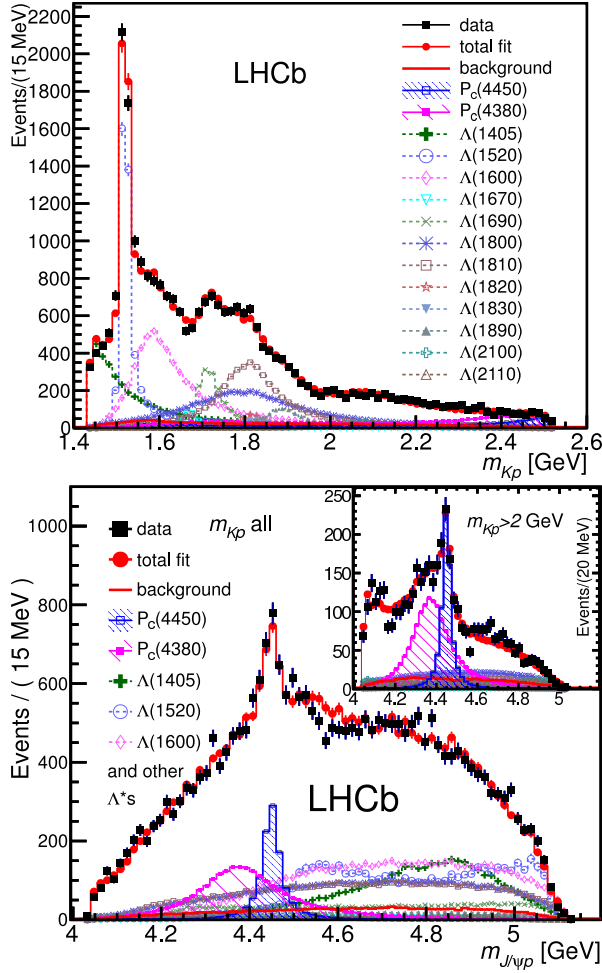
**Figure 1:** Dalitz plot distributions for  $\Lambda_b^0 \rightarrow J\psi p K^-$  decays as observed by LHCb.

An amplitude analysis was performed to clarify the nature of this band that followed in the footsteps of a similar analysis of  $\bar{B}^0 \rightarrow \psi(2S)\pi^+K^-$  ( $\psi(2S) \rightarrow \mu^+\mu^-$ ) performed by the LHCb a year earlier in which the  $Z(4430)^+$  tetraquark candidate [9] was confirmed and the resonant character of its amplitude was demonstrated by an Argand diagram [10]. The final states are very similar, with  $\pi^+$  being replaced by  $p$ . The signal statistics,  $26\,000 \pm 166$ , and the background level, 5.4%, are also very comparable. The quasi-two-body amplitude model was constructed based on an isobar approximation (*i.e.* summing up Breit-Wigner amplitudes) and helicity formalism to parameterize dynamics of contributing decay processes. The amplitude fit spanned a kinematically complete, six-dimensional space of independent kinematic variables. All six dimensions of  $\Lambda_b$  decay kinematics were used in the amplitude fit, including invariant masses of  $K^-p$  ( $m_{Kp}$ ) and  $J\psi p$ , ( $m_{J\psi p}$ ) helicity angles ( $\theta$ ) of  $\Lambda_b$ ,  $J\psi$ ,  $\Lambda^*$  or pentaquark candidate  $P_c^+ \rightarrow J\psi p$ , and angles between decay planes of the particles. Fourteen reasonably well established  $\Lambda^*$  resonances were considered with masses and widths fixed to the values listed in 2014 PDG edition [11], and varied within their uncertainties when evaluating systematic errors. Their helicity couplings (1-6 complex numbers per resonance) were determined from the fit to the data. It was found that the  $\Lambda^*$  contributions alone failed to describe the data and it was necessary to add two exotic  $P_c^+ \rightarrow J\psi p$  contributions to the matrix element (10 free parameters per resonance), before the narrow structure seen in  $m_{J\psi p}$  could be reasonably well reproduced, as illustrated in Fig. 2.

The lower mass state,  $P_c(4380)^+$ , has a fitted mass of  $4380 \pm 8 \pm 29$  MeV, width of  $205 \pm 18 \pm 86$  MeV, fit fraction of  $8.4 \pm 0.7 \pm 4.2$  % and significance of  $9\sigma$ . The higher mass state,  $P_c(4450)^+$ , has a fitted mass of  $4449.8 \pm 1.7 \pm 2.5$  MeV, narrower width of  $39 \pm 5 \pm 19$  MeV, a fit fraction of  $4.1 \pm 0.5 \pm 1.1$  % and significance of  $12\sigma$ . The need for a second  $P_c^+$  state becomes visually apparent in the  $m_{J\psi p}$  distribution for events with high values of  $m_{Kp}$ , where  $\Lambda^*$  contributions are the smallest (in the inset of Fig. 2). Even though contributions from the two  $P_c^+$  states are most visible in this region, they interfere destructively in this part of the Dalitz plane. The constructive  $P_c^+$  interference makes their combined contribution the largest at the other end of their band on the Dalitz plane, corresponding to the opposite end of the  $\cos\theta_{P_c^+}$  distribution (see Fig. 8b in Ref. 8). This pattern requires them to be of opposite parity. A similar interference pattern is observed in the  $\cos\theta_{\Lambda^*}$  distribution (Fig. 7 in Ref. 8), which is a consequence of parity-doublets in the  $\Lambda^*$  spectrum. Unfortunately, spins of the two  $P_c^+$  states were not uniquely determined. Within the statistical and systematic ambiguities,  $(3/2, 5/2)$  and  $(5/2, 3/2)$  combinations with either  $(-, +)$  or  $(+, -)$  parities, were not well resolved. The other combinations were disfavored. The Argand diagrams for the two  $P_c^+$  states are shown in Fig. 3. They were obtained by replacing the Breit-Wigner amplitude for one of the  $P_c^+$  states at a time by a combination of independent complex amplitudes at six equidistant points in the  $\pm\Gamma_0$  range

# Baryon Particle Listings

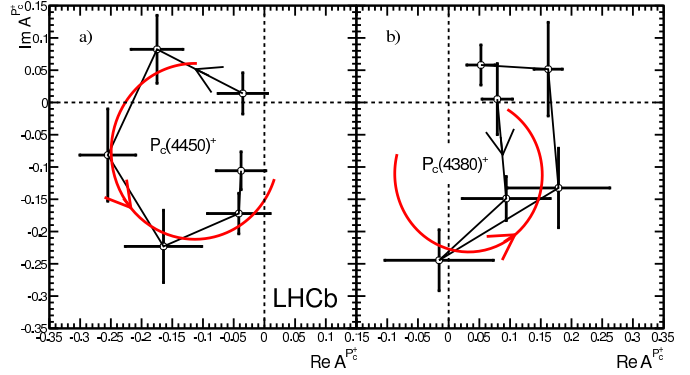
## Pentaquarks



**Figure 2:** Projections of the amplitude fits with  $P_c(4380)^+$  and  $P_c(4450)^+$  states to the  $\Lambda_b^0 \rightarrow J\psi p K^-$  data onto the invariant mass distributions of  $m_{Kp}$  (left) and  $m_{J\psi p}$  (right).

(interpolated in mass for continuity) which were fit to the data simultaneously with the other parameters of the full matrix element model. While the narrower  $P_c(4450)^+$  state shows the expected resonant behavior, the diagram for  $P_c(4380)^+$  deviates somewhat from the expectation. The statistical errors are large, especially for the broader  $P_c(4380)^+$  state. Higher statistics data might make these diagrams more conclusive. The addition of further  $\Lambda$  states beyond the well-established ones, of  $\Sigma$  excitations (expected to be suppressed) and of non-resonant contributions with a constant amplitude, did not remove the need for two pentaquark states in the model to describe the data. Yet  $\Lambda^*$  spectroscopy is a complex problem, from both experimental and theoretical points of view. This is illustrated by the recent reanalysis of  $\bar{K}N$  scattering data [12] in which the  $\Lambda(1800)$  state, which was previously considered to be “well established”, is not seen, and where evidence for a few previously unidentified states is included. In fact, all theoretical models of  $\Lambda^*$  baryons [13–18] predict a much larger number of higher mass excitations than is established experimentally. Because of the

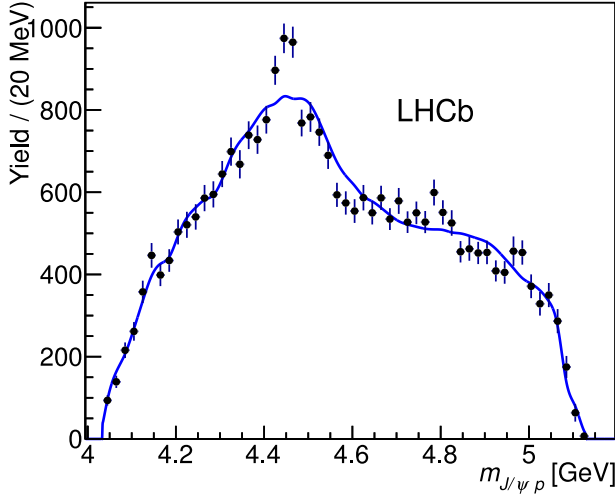
high density of predicted states, presumably with large widths, these may be difficult to identify experimentally. Non-resonant contributions with a non-trivial  $K^-p$  mass dependence may also be present. Therefore, LHCb also inspected their data with an approach that is nearly model-independent with respect to  $K^-p$  contributions [19].



**Figure 3:** Fitted values of the real and imaginary parts of the amplitudes of the  $P_c(4450)^+$  (left) and  $P_c(4380)^+$  (right) states for  $\Lambda_b^0 \rightarrow J\psi p K^-$  shown in the Argand diagrams as connected points with the error bars (masses increase counterclockwise). The solid red curves are the predictions from the Breit-Wigner formula, with resonance masses and widths set to the nominal fit results, scaled to the displayed points.

A representation of the Dalitz plane distribution was constructed using the observed  $m_{Kp}$  distribution and Legendre polynomial moments of the cosine of the  $\Lambda^*$  helicity angle determined from the data as a function of  $m_{Kp}$ . The maximal rank of the moments generated by the  $K^-p$  contributions alone cannot be higher than twice the largest total angular momentum. Since high-spin  $\Lambda^*$  states cannot significantly contribute at low  $m_{Kp}$  values, high rank moments were excluded from the representation (see Fig. 1 and 3 in Ref. 19). When projected onto  $m_{J\psi p}$  axis of the Dalitz plane, this representation cannot describe the data as shown in Fig. 4. The disagreement was quantified to be at least  $9\sigma$ , thus the hypothesis that only  $K^-p$  contributions can generate the observed  $m_{J\psi p}$  mass structure could be rejected with very high confidence without any assumptions about number of  $K^-p$  contributions, their resonant or non-resonant character, their mass shapes or their interference patterns. This proved a need for contributions from exotic hadrons or from rescattering effects of conventional ones. However, this approach is not suitable for their characterization.

Many theoretical groups interpreted the  $P_c^+$  states in terms of diquarks and triquarks as building blocks of a compact pentaquark [20–26]. The pair of states of opposite parity with the  $3/2$  spin assignment to  $P_c(4380)^+$  and  $5/2$  to  $P_c(4450)^+$



**Figure 4:** The efficiency-corrected and background-subtracted distribution of  $m_{J/\psi p}$  for the data (black points with error bars), with the reflection of  $K^-p$  mass distribution and of the moments of the  $K^-p$  helicity angle, which can be accommodated by any plausible  $K^-p$  contribution (solid blue line) superimposed. The data and the reflection are inconsistent at  $> 9\sigma$  level.

can be achieved by increasing the angular momentum between the constituents by one unit, which can also make the heavier state narrower. However, their mass splitting is too small to be only due to this mechanism [20] and requires fine-tuning of such models. It is also not clear if centrifugal barrier factor provides enough width suppression via spatial separation of  $c$  and  $\bar{c}$  quarks to explain the width ratio between the two  $P_c^+$  states and the narrowness of  $P_c(4450)^+$  in absolute units as the phase space for  $J/\psi p$  decay is very large (more than 400 MeV).

More effective width suppression mechanism is offered by a loosely bound charmed baryon-anticharmed meson molecular model, in which  $c$  and  $\bar{c}$  can be separated to much larger distances resulting in a smaller probability of them getting close to each other in order to make a  $J/\psi$ . Since molecular binding energy cannot be large, masses of such molecules must be near the sum of the baryon and meson masses. The narrowness of  $P_c(4450)^+$  and its proximity to appropriate baryon-meson mass threshold make the molecular model attractive in spite of its inability to account for other features of the LHCb results (see below).

In order to view the narrow pentaquark in a wider perspective, it is useful to consider it together with several analogous exotic states with hidden charm and bottom in the meson sector. This provides additional significant motivation for the molecular model. At least five exotic mesons are close to thresholds of two heavy-light mesons:  $X(3872)$  [27–30],  $Z_b(10610)$  and  $Z_b(10650)$  in the bottomonium sector [31–35]

and  $Z_c(3900)$  [36–40] and  $Z_c(4020/4025)$  [41–43] in the charmonium sector (see Table II of Ref. 44). They share several important features: a) their masses are near thresholds and their spin and parity correspond to  $S$ -wave combination of the two mesons; b) they are very narrow, despite very large phase space for decay into quarkonium + pion(s); c) the branching fractions for “fall apart” mode into two mesons are much larger than branching fractions for decay into quarkonium and pion(s); d) there are no states at two pseudoscalar thresholds ( $\bar{D}D$  and  $\bar{B}B$ ), where there can be no binding through pseudoscalar exchange.

The above provide a strong hint that these states are deuteron-like loosely bound states of two heavy mesons [45–53]. It is then natural to conjecture that similar bound states might exist of two heavy baryons [54,55], or a meson and a baryon or a baryon and an antibaryon, leading to a rather accurate prediction of the  $P_c(4450)^+$  mass as  $3/2^- \Sigma_c \bar{D}^*$  molecule: 4462.4 MeV [56,44]. It is essential that the two hadrons be heavy, in order to minimize the repulsive kinetic energy [54–57].

One may also consider a wider framework of doubly heavy baryon-meson hadronic molecules, which might include mixtures of various two-hadron states [58,59]. In this context it is important to keep in mind that the molecule’s width cannot be smaller than the sum of its constituents’ widths [60–62].

Following the LHCb discovery, several groups carried out a detailed analysis of the  $P_c^+$  states as hadronic molecules [63–71]. The molecular picture has also been extended to a hadronic molecule built from a colored “baryon” and “meson” [72].

When trying to interpret both  $P_c(4380)^+$  and  $P_c(4380)^+$  as hadronic molecules, it is essential to remember that these two states have opposite parities. Thus one cannot construct both of them as  $S$ -wave bound states of a meson and a baryon with natural parities. Therefore, the interpretation of the  $P_c^+$  states as hadronic molecules has been by no means unanimous. Moreover, the molecular model is not consistent with one of the  $P_c^+$  states having a spin of  $5/2$ , since  $S$ -wave combinations of baryon-meson combination that can produce such spin have thresholds which are too high in mass to be plausible. Therefore, the confirmation or disproof of the presence of this high-spin structure is a critical test of the molecular model. The large  $P_c(4380)^+$  width is also difficult to accommodate in the molecular bound state model, but could have its origin in baryon-meson rescattering effects discussed below.

Shortly after the experimental discovery it has been conjectured that the observed resonances could be kinematic effects due to vicinity of thresholds and so-called triangle singularity [73–76]. While these effects might explain the large  $P_c(4380)^+$  width, since such models involve  $S$ -wave rescattering of virtual baryon-meson pairs, they also cannot be reconciled with one of the  $P_c^+$  peaks having effective spin of  $5/2$ .

In addition to the molecular and diquark approach, the  $P_c^+$  pentaquarks have also been analysed within the soliton

# Baryon Particle Listings

## Pentaquarks

picture of baryons, as a bound state of a soliton and an anti-charmed meson [77]. Quite recently an interesting attempt has been made to explain the narrow width of tetraquarks and pentaquarks by extending to these states the string junction picture of baryons in QCD [78].

More extensive reviews of the theoretical issues can be found in Refs. 79,80.

So far the  $P_c^+$  states have been observed by only one experiment in only one channel. It is essential to explore other possible experimental channels. Proposals have been made for searching for heavy pentaquarks in photoproduction [81–83], (c.f. also related work on computation of  $J/\psi(\eta_c)N$  and  $\Upsilon(\eta_b)N$  cross sections [84]), in heavy ion collisions at LHC [85], in  $pA$  collisions [86], and in pion-induced processes [87,88].

### References

1. T.G. Trippe *et al.* (Particle Data Group), Rev. Mod. Phys. **48**, S1 (1976), [Erratum: Rev. Mod. Phys. **48**, 497 (1976)].
2. K. Hikasa *et al.* (Particle Data Group), Phys. Rev. **D45**, S1 (1992), [Erratum: Phys. Rev. **D46**, 5210 (1992)].
3. M. Praszalowicz, *Skyrmions and Anomalies*, p.112, M. Jezabek Ed., World Scientific Publishing(1987), ISBN 9971503506.
4. D. Diakonov, V. Petrov and M.V. Polyakov, Z. Phys. **A359**, 305 (1997) [hep-ph/9703373].
5. H. Weigel, Eur. Phys. J. **A2**, 391 (1998) [hep-ph/9804260].
6. K.H. Hicks, Eur. Phys. J. **H37**, 1 (2012).
7. W.M. Yao *et al.* (Particle Data Group), J. Phys. **G33**, 1 (2006).
8. R. Aaij *et al.* (LHCb), Phys. Rev. Lett. **115**, 072001 (2015) [arXiv:1507.03414].
9. S. Choi *et al.* (Belle), Phys. Rev. Lett. **100**, 142001 (2008) [arXiv:0708.1790].
10. R. Aaij *et al.* (LHCb), Phys. Rev. Lett. **112**, 222002 (2014) [arXiv:1404.1903].
11. K. Olive *et al.* (Particle Data Group), Chin. Phys. C **38**, 090001 (2014) [arXiv:1412.1408].
12. C. Fernandez-Ramirez *et al.*, (2015), [arXiv:1510.07065].
13. R. Faustov and V. Galkin, Phys. Rev. **D92**, 054005 (2015) [arXiv:1507.04530].
14. S. Capstick and N. Isgur, Phys. Rev. **D34**, 2809 (1986).
15. U. Loring, B. Metsch and H. Petry, Eur. Phys. J. **A10**, 447 (2001).
16. T. Melde, W. Plessas and B. Sengl, Phys. Rev. **D77**, 114002 (2008).
17. E. Santopinto and J. Ferretti, Phys. Rev. **C92**, 025202 (2015) [arXiv:1412.7571].
18. G. Engel *et al.*, Phys. Rev. **D87**, 074504 (2013) [arXiv:1301.4318].
19. R. Aaij *et al.* (LHCb) (2016), arXiv:1604.05708.
20. L. Maiani, A.D. Polosa and V. Riquer, Phys. Lett. **B749**, 289 (2015) [arXiv:1507.04980].
21. R.F. Lebed, Phys. Lett. **B749**, 454 (2015) [arXiv:1507.05867].
22. V.V. Anisovich *et al.*, (2015) [arXiv:1507.07652].
23. G.-N. Li, X.-G. He and M. He, JHEP **12**, 128 (2015) [arXiv:1507.08252].
24. R. Ghosh, A. Bhattacharya and B. Chakrabarti (2015), arXiv:1508.00356.
25. Z.-G. Wang, Eur. Phys. J. **C76**, 70 (2016) [arXiv:1508.01468].
26. R. Zhu and C.-F. Qiao, Phys. Lett. **B756**, 259 (2016) [arXiv:1510.08693].
27. S.K. Choi *et al.* (Belle), Phys. Rev. Lett. **91**, 262001 (2003) [hep-ex/0309032].
28. D. Acosta *et al.* (CDF), Phys. Rev. Lett. **93**, 072001 (2004) [hep-ex/0312021].
29. B. Aubert *et al.* (BaBar), Phys. Rev. **D71**, 071103 (2005) [hep-ex/0406022].
30. V. M. Abazov *et al.* (D0), Phys. Rev. Lett. **93**, 162002 (2004) [hep-ex/0405004].
31. M. Karliner and H. J. Lipkin (2008) [arXiv:0802.0649].
32. K.F. Chen *et al.* (Belle), Phys. Rev. Lett. **100**, 112001 (2008) [arXiv:0710.2577].
33. A. Bondar *et al.* (Belle), Phys. Rev. Lett. **108**, 122001 (2012) [arXiv:1110.2251].
34. P. Krokovny *et al.* (Belle), Phys. Rev. **D88**, 052016 (2013) [arXiv:1308.2646].
35. A. Garmash *et al.* (Belle), Phys. Rev. **D91**, 072003 (2015) [arXiv:1403.0992].
36. M. Ablikim *et al.* (BES III), Phys. Rev. Lett. **110**, 252001 (2013) [arXiv:1303.5949].
37. Z. Q. Liu *et al.* (Belle), Phys. Rev. Lett. **110**, 252002 (2013) [arXiv:1304.0121].
38. T. Xiao *et al.*, Phys. Lett. **B727**, 366 (2013) [arXiv:1304.3036].
39. M. Ablikim *et al.* (BES III), Phys. Rev. Lett. **112**, 022001 (2014) [arXiv:1310.1163].
40. M. Ablikim *et al.* (BES III), Phys. Rev. Lett. **115**, 112003 (2015) [arXiv:1506.06018].
41. M. Ablikim *et al.* (BES III), Phys. Rev. Lett. **111**, 242001 (2013) [arXiv:1309.1896].
42. M. Ablikim *et al.* (BES III), Phys. Rev. Lett. **113**, 212002 (2014) [arXiv:1409.6577].
43. M. Ablikim *et al.* (BES III), Phys. Rev. Lett. **112**, 132001 (2014) [arXiv:1308.2760].
44. M. Karliner, Acta Phys. Polon. **B47**, 117 (2016).
45. M.B. Voloshin and L. B. Okun, Sov. Phys. JETP Lett. **23**, 333 (1976) [Pisma Zh. Eksp. Teor. Fiz. **23**, 369 (1976)].
46. A. De Rujula, H. Georgi and S. Glashow, Phys. Rev. Lett. **38**, 317 (1977).
47. N.A. Tornqvist, Phys. Rev. Lett. **67**, 556 (1991).
48. N.A. Tornqvist, Z. Phys. **C61**, 525 (1994) [hep-ph/9310247].
49. N.A. Tornqvist, Phys. Lett. **B590**, 209 (2004) [hep-ph/0402237].
50. C.E. Thomas and F.E. Close, Phys. Rev. **D78**, 034007 (2008) [arXiv:0805.3653].
51. M. Suzuki, Phys. Rev. **D72**, 114013 (2005) [hep-ph/0508258].
52. S. Fleming *et al.*, Phys. Rev. **D76**, 034006 (2007) [hep-ph/0703168].
53. T.E.O. Ericson and G. Karl, Phys. Lett. **B309**, 426 (1993).

# Baryon Particle Listings

## Pentaquarks, $P_c(4380)^+$ , $P_c(4450)^+$

54. M. Karliner, H.J. Lipkin and N.A. Tornqvist, in *Proceedings, 14th International Conference on Hadron spectroscopy (Hadron 2011)*, (2011) [arXiv:1109.3472].
55. M. Karliner, H.J. Lipkin and N.A. Tornqvist, Nucl. Phys. (Proc. Supp.) **102**, 225 (2012).
56. M. Karliner and J.L. Rosner, Phys. Rev. Lett. **115**, 122001 (2015) [arXiv:1506.06386].
57. X.-Q. Li and X. Liu, Eur. Phys. J. **C74**, 3198 (2014) [arXiv:1409.3332].
58. J.-J. Wu *et al.*, Phys. Rev. Lett. **105**, 232001 (2010) [arXiv:1007.0573].
59. Z.-C. Yang *et al.*, Chin. Phys. C **36**, 6 (2012) [arXiv:1105.2901].
60. C. Hanhart, Yu.S. Kalashnikova and A.V. Nefediev, Phys. Rev. **D81**, 094028 (2010) [arXiv:1002.4097].
61. A.A. Filin *et al.*, Phys. Rev. Lett. **105**, 019101 (2010) [arXiv:1004.4789].
62. F.-K. Guo and U.-G. Meissner, Phys. Rev. **D84**, 014013 (2011) [arXiv:1102.3536].
63. R. Chen *et al.*, Phys. Rev. Lett. **115**, 132002 (2015) [arXiv:1507.03704].
64. H.-X. Chen *et al.*, Phys. Rev. Lett. **115**, 172001 (2015) [arXiv:1507.03717].
65. L. Roca, J. Nieves and E. Oset, Phys. Rev. **D92**, 094003 (2015) [arXiv:1507.04249].
66. J. He, Phys. Lett. **B753**, 547 (2016) [arXiv:1507.05200].
67. H. Huang *et al.*, (2015), arXiv:1510.04648.
68. L. Roca and E. Oset (2016), arXiv:1602.06791.
69. Q.-F. Lu and Y.-B. Dong (2016), arXiv:1603.00559.
70. Y. Shimizu, D. Suenaga and M. Harada (2016), arXiv:1603.02376.
71. C.-W. Shen *et al.*, (2016), arXiv:1603.04672.
72. A. Mironov and A. Morozov, Sov. Phys. JETP Lett. **102**, 271 (2015) [arXiv:1507.04694].
73. F.-K. Guo *et al.*, Phys. Rev. **D92**, 071502 (2015) [arXiv:1507.04950].
74. U.-G. Meissner and J. A. Oller, Phys. Lett. **B751**, 59 (2015) [arXiv:1507.07478].
75. X.-H. Liu, Q. Wang and Q. Zhao (2015), arXiv:1507.05359.
76. M. Mikhasenko (2015), arXiv:1507.06552.
77. N.N. Scoccola, D.O. Riska and M. Rho, Phys. Rev. **D92**, 051501 (2015) [arXiv:1508.01172].
78. G. Rossi and G. Veneziano (2016), arXiv:1603.05830.
79. T.J. Burns, Eur. Phys. J. **A51**, 152 (2015) [arXiv:1509.02460].
80. H.-X. Chen *et al.*, (2016), arXiv:1601.02092.
81. Q. Wang, X.-H. Liu and Q. Zhao, Phys. Rev. **D92**, 034022 (2015) [arXiv:1508.00339].
82. V. Kubarovsky and M.B. Voloshin, Phys. Rev. **D92**, 031502 (2015) [arXiv:1508.00888].
83. M. Karliner and J.L. Rosner, Phys. Lett. **B752**, 329 (2016) [arXiv:1508.01496].
84. C.W. Xiao and U.-G. Meissner, Phys. Rev. **D92**, 114002 (2015) [arXiv:1508.00924].
85. R.-Q. Wang *et al.*, (2016), arXiv:1601.02835.
86. I. Schmidt and M. Siddikov (2016), arXiv:1601.05621.
87. Q.-F. Lu *et al.*, Phys. Rev. **D93**, 034009 (2016) [arXiv:1510.06271].
88. X.-H. Liu and M. Oka (2016), arXiv:1602.07069.

### $P_c(4380)^+$

Status: \*

A resonance seen in  $\Lambda_b^0 \rightarrow P_c^+ K^-$ , then  $P_c \rightarrow J/\psi p$ , with a significance of 9 standard deviations. The  $J/\psi p$  quark content is  $uudc\bar{c}$ , a pentaquark. See also the  $P_c(4450)^+$ . In the best amplitude fit, the two states have opposite parity, one having  $J = 3/2$ , the other  $J = 5/2$ .

#### $P_c(4380)^+$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4380 ± 8 ± 29</b>	AAIJ	15P	LHCB <i>pp</i> at 7, 8 TeV

#### $P_c(4380)^+$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>205 ± 18 ± 86</b>	AAIJ	15P	LHCB <i>pp</i> at 7, 8 TeV

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $J/\psi p$	seen

#### $P_c(4380)^+$ BRANCHING RATIOS

$\Gamma(J/\psi p)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AAIJ	15P	LHCB	$pp$ at 7, 8 TeV

AAIJ 15P PRL 115 072001 R. Aaij *et al.* (LHCb Collab.)

### $P_c(4450)^+$

Status: \*

A resonance seen in  $\Lambda_b^0 \rightarrow P_c^+ K^-$ , then  $P_c \rightarrow J/\psi p$ , with a significance of 12 standard deviations. The  $J/\psi p$  quark content is  $uudc\bar{c}$ , a pentaquark. See also the  $P_c(4380)^+$ . In the best amplitude fit, the two states have opposite parity, one having  $J = 3/2$ , the other  $J = 5/2$ .

#### $P_c(4450)^+$ MASS

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4449.8 ± 1.7 ± 2.5</b>	AAIJ	15P	LHCB <i>pp</i> at 7, 8 TeV

#### $P_c(4450)^+$ WIDTH

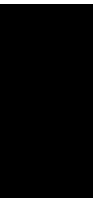
VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>39 ± 5 ± 19</b>	AAIJ	15P	LHCB <i>pp</i> at 7, 8 TeV

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $J/\psi p$	seen

#### $P_c(4450)^+$ BRANCHING RATIOS

$\Gamma(J/\psi p)/\Gamma_{\text{total}}$				$\Gamma_1/\Gamma$
VALUE	DOCUMENT ID	TECN	COMMENT	
seen	AAIJ	15P	LHCB	$pp$ at 7, 8 TeV

AAIJ 15P PRL 115 072001 R. Aaij *et al.* (LHCb Collab.)





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# SEARCHES FOR MONOPOLES, SUPERSYMMETRY, TECHNICOLOR, COMPOSITENESS, EXTRA DIMENSIONS, etc.

## Magnetic Monopole Searches

### MAGNETIC MONOPOLES

Updated August 2015 by D. Milstead (Stockholm Univ.) and E.J. Weinberg (Columbia Univ.).

The symmetry between electric and magnetic fields in the sourcefree Maxwell's equations naturally suggests that electric charges might have magnetic counterparts, known as magnetic monopoles. Although the greatest interest has been in the supermassive monopoles that are a firm prediction of all grand unified theories, one cannot exclude the possibility of lighter monopoles, even though there is at present no strong theoretical motivation for these.

In either case, the magnetic charge is constrained by a quantization condition first found by Dirac [1]. Consider a monopole with magnetic charge  $Q_M$  and a Coulomb magnetic field

$$\mathbf{B} = \frac{Q_M}{4\pi} \frac{\hat{\mathbf{r}}}{r^2}. \quad (1)$$

Any vector potential  $\mathbf{A}$  whose curl is equal to  $\mathbf{B}$  must be singular along some line running from the origin to spatial infinity. This Dirac string singularity could potentially be detected through the extra phase that the wavefunction of a particle with electric charge  $Q_E$  would acquire if it moved along a loop encircling the string. For the string to be unobservable, this phase must be a multiple of  $2\pi$ . Requiring that this be the case for any pair of electric and magnetic charges gives the condition that all charges be integer multiples of minimum charges  $Q_E^{\min}$  and  $Q_M^{\min}$  obeying

$$Q_E^{\min} Q_M^{\min} = 2\pi. \quad (2)$$

(For monopoles which also carry an electric charge, called dyons, the quantization conditions on their electric charges can be modified. However, the constraints on magnetic charges, as well as those on all purely electric particles, will be unchanged.)

Another way to understand this result is to note that the conserved orbital angular momentum of a point electric charge moving in the field of a magnetic monopole has an additional component, with

$$\mathbf{L} = m\mathbf{r} \times \mathbf{v} - 4\pi Q_E Q_M \hat{\mathbf{r}} \quad (3)$$

Requiring the radial component of  $\mathbf{L}$  to be quantized in half-integer units yields Eq. (2).

If there are unbroken gauge symmetries in addition to the U(1) of electromagnetism, the above analysis must be modified [2,3]. For example, a monopole could have both a

U(1) magnetic charge and a color magnetic charge. The latter could combine with the color charge of a quark to give an additional contribution to the phase factor associated with a loop around the Dirac string, so that the U(1) charge could be the Dirac charge  $Q_M^D \equiv 2\pi/e$ , the result that would be obtained by substituting the electron charge into Eq. (2). On the other hand, for monopoles without color-magnetic charge, one would simply insert the quark electric charges into Eq. (2) and conclude that  $Q_M$  must be a multiple of  $6\pi/e$ .

The prediction of GUT monopoles arises from the work of 't Hooft [4] and Polyakov [5], who showed that certain spontaneously broken gauge theories have nonsingular classical solutions that lead to magnetic monopoles in the quantum theory. The simplest example occurs in a theory where the vacuum expectation value of a triplet Higgs field  $\phi$  breaks an SU(2) gauge symmetry down to the U(1) of electromagnetism and gives a mass  $M_V$  to two of the gauge bosons. In order to have finite energy,  $\phi$  must approach a vacuum value at infinity. However, there is a continuous family of possible vacua, since the scalar field potential determines only the magnitude  $v$  of  $\langle\phi\rangle$ , but not its orientation in the internal SU(2) space. In the monopole solution, the direction of  $\phi$  in internal space is correlated with the position in physical space; *i.e.*,  $\phi^a \sim v\hat{r}^a$ . The stability of the solution follows from the fact that this twisting Higgs field cannot be smoothly deformed to a spatially uniform vacuum configuration. Reducing the energetic cost of the spatial variation of  $\phi$  requires a nonzero gauge potential, which turns out to yield the magnetic field corresponding to a charge  $Q_M = 4\pi/e$ . Numerical solution of the classical field equations shows that the mass of this monopole is

$$M_{\text{mon}} \sim \frac{4\pi M_V}{e^2}. \quad (4)$$

The essential ingredient here was the fact that the Higgs fields at spatial infinity could be arranged in a topologically nontrivial configuration. A discussion of the general conditions under which this is possible is beyond the scope of this review, so we restrict ourselves to the two phenomenologically most important cases.

The first is the electroweak theory, with SU(2) × U(1) broken to U(1). There are no topologically nontrivial configurations of the Higgs field, and hence no topologically stable monopole solutions.

The second is when any simple Lie group is broken to a subgroup with a U(1) factor, a case that includes all grand unified theories. The monopole mass is determined by the mass scale of the symmetry breaking that allows nontrivial topology. For example, an SU(5) model with

$$\text{SU}(5) \xrightarrow{M_X} \text{SU}(3) \times \text{SU}(2) \times \text{U}(1) \xrightarrow{M_W} \text{SU}(3) \times \text{U}(1) \quad (5)$$

has a monopole [6] with  $Q_M = 2\pi/e$  and mass

$$M_{\text{mon}} \sim \frac{4\pi M_X}{g^2}, \quad (6)$$

## Searches Particle Listings

### Magnetic Monopole Searches

where  $g$  is the SU(5) gauge coupling. For a unification scale of  $10^{16}$  GeV, these monopoles would have a mass  $M_{\text{mon}} \sim 10^{17} - 10^{18}$  GeV.

In theories with several stages of symmetry breaking, monopoles of different mass scales can arise. In an SO(10) theory with

$$\text{SO}(10) \xrightarrow{M_1} \text{SU}(4) \times \text{SU}(2) \times \text{SU}(2) \xrightarrow{M_2} \text{SU}(3) \times \text{SU}(2) \times \text{U}(1) \quad (7)$$

there is monopole with  $Q_M = 2\pi/e$  and mass  $\sim 4\pi M_1/g^2$  and a much lighter monopole with  $Q_M = 4\pi/e$  and mass  $\sim 4\pi M_2/g^2$  [7].

The central core of a GUT monopole contains the fields of the superheavy gauge bosons that mediate baryon number violation, so one might expect that baryon number conservation could be violated in baryon-monopole scattering. The surprising feature, pointed out by Callan [8] and Rubakov [9], is that these processes are not suppressed by powers of the gauge boson mass. Instead, the cross-sections for catalysis processes such as  $p + \text{monopole} \rightarrow e^+ + \pi^0 + \text{monopole}$  are essentially geometric; *i.e.*,  $\sigma_{\Delta B} \sim 10^{-27} \text{ cm}^2$ , where  $\beta = v/c$ . Note, however, that intermediate mass monopoles arising at later stages of symmetry breakings, such as the doubly charged monopoles of the SO(10) theory, do not catalyze baryon number violation.

**Production and Annihilation:** GUT monopoles are far too massive to be produced in any foreseeable accelerator. However, they could have been produced in the early universe as topological defects arising via the Kibble mechanism [10] in a symmetry-breaking phase transition. Estimates of the initial monopole abundance, and of the degree to which it can be reduced by monopole-antimonopole annihilation, predict a present-day monopole abundance that exceeds by many orders of magnitude the astrophysical and experimental bounds described below [11]. Cosmological inflation and other proposed solutions to this primordial monopole problem generically lead to present-day abundances exponentially smaller than could be plausibly detected, although potentially observable abundances can be obtained in scenarios with carefully tuned parameters.

If monopoles light enough to be produced at colliders exist, one would expect that these could be produced by analogs of the electromagnetic processes that produce pairs of electrically charged particles. Because of the large size of the magnetic charge, this is a strong coupling problem for which perturbation theory cannot be trusted. Indeed, the problem of obtaining reliable quantitative estimates of the production cross-sections remains an open one, on which there is no clear consensus.

**Astrophysical and Cosmological Bounds:** If there were no galactic magnetic field, one would expect monopoles in the galaxy to have typical velocities of the order of  $10^{-3}c$ , comparable to the virial velocity in the galaxy (relevant if the monopoles cluster with the galaxy) and the peculiar velocity of the galaxy with respect to the CMB rest frame (relevant if the monopoles are not bound to the galaxy). This situation is modified by the existence of a galactic magnetic field  $B \sim 3\mu\text{G}$ .

A monopole with the Dirac charge and mass  $M$  would be accelerated by this field to a velocity

$$v_{\text{mag}} \sim \begin{cases} c, & M \lesssim 10^{11} \text{ GeV} \\ 10^{-3}c \left( \frac{10^{17} \text{ GeV}}{M} \right)^{1/2}, & M \gtrsim 10^{11} \text{ GeV} \end{cases} \quad (8)$$

Accelerating these monopoles drains energy from the magnetic field. Parker [12] obtained an upper bound on the flux of monopoles in the galaxy by requiring that the rate of this energy loss be small compared to the time scale on which the galactic field can be regenerated. With reasonable choices for the astrophysical parameters (see Ref. 13 for details), this Parker bound is

$$F < \begin{cases} 10^{-15} \text{ cm}^{-2} \text{ sr}^{-1} \text{ sec}^{-1}, & M \lesssim 10^{17} \text{ GeV} \\ 10^{-15} \left( \frac{M}{10^{17} \text{ GeV}} \right) \text{ cm}^{-2} \text{ sr}^{-1} \text{ sec}^{-1}, & M \gtrsim 10^{17} \text{ GeV} \end{cases} \quad (9)$$

Applying similar arguments to an earlier seed field that was the progenitor of the current galactic field leads to a tighter bound [14],

$$F < \left[ \frac{M}{10^{17} \text{ GeV}} + (3 \times 10^{-6}) \right] 10^{-16} \text{ cm}^{-2} \text{ sr}^{-1} \text{ sec}^{-1}. \quad (10)$$

Considering magnetic fields in galactic clusters gives a bound [15] which, although less secure, is about three orders of magnitude lower than the Parker bound.

A flux bound can also be inferred from the total mass of monopoles in the universe. If the monopole mass density is a fraction  $\Omega_M$  of the critical density, and the monopoles were uniformly distributed throughout the universe, there would be a monopole flux

$$F_{\text{uniform}} = 1.3 \times 10^{-16} \Omega_M \left( \frac{10^{17} \text{ GeV}}{M} \right) \left( \frac{v}{10^{-3}c} \right) \text{ cm}^{-2} \text{ sr}^{-1} \text{ sec}^{-1}. \quad (11)$$

If we assume that  $\Omega_M \sim 0.1$ , this gives a stronger constraint than the Parker bound for  $M \sim 10^{15}$  GeV. However, monopoles with masses  $\sim 10^{17}$  GeV are not ejected by the galactic field and can be gravitationally bound to the galaxy. In this case their flux within the galaxy is increased by about five orders of magnitude for a given value of  $\Omega_M$ , and the mass density bound only becomes stronger than the Parker bound for  $M \sim 10^{18}$  GeV.

A much more stringent flux bound applies to GUT monopoles that catalyze baryon number violation. The essential idea is that compact astrophysical objects would capture monopoles at a rate proportional to the galactic flux. These monopoles would then catalyze proton decay, with the energy released in the decay leading to an observable increase in the luminosity of the object. A variety of bounds, based on neutron stars [16–20], white dwarfs [21], and Jovian planets [22] have been obtained. These depend in the obvious manner on the catalysis cross section, but also on the details of the astrophysical scenarios; *e.g.*, on how much the accumulated density is reduced by monopole-antimonopole annihilation, and

on whether monopoles accumulated in the progenitor star survive its collapse to a white dwarf or neutron star. The bounds obtained in this manner lie in the range

$$F\left(\frac{\sigma_{\Delta B}\beta}{10^{-27}\text{cm}^2}\right) \sim (10^{-18} - 10^{-29})\text{cm}^{-2}\text{sr}^{-1}\text{sec}^{-1}. \quad (12)$$

It is important to remember that not all GUT monopoles catalyze baryon number nonconservation. In particular, the intermediate mass monopoles that arise in some GUTs at later stages of symmetry-breaking are examples of theoretically motivated monopoles that are exempt from the bound of Eq. (12).

**Searches for Magnetic Monopoles:** To date there have been no confirmed observations of exotic particles possessing magnetic charge. Precision measurements of the properties of known particles have led to tight limits on the values of magnetic charge they may possess. Using the induction method (see below), the electron's magnetic charge has been found to be  $Q_e^m < 10^{-24}Q_M^D$  [23] (where  $Q_M^D$  is the Dirac charge). Furthermore, measurements of the anomalous magnetic moment of the muon have been used to place a model dependent lower limit of 120 GeV on the monopole mass <sup>1</sup> [24]. Nevertheless, guided mainly by Dirac's argument and the predicted existence of monopoles from spontaneous symmetry breaking mechanisms, searches have been routinely made for monopoles produced at accelerators, in cosmic rays, and bound in matter [25]. Although the resultant limits from such searches are usually made under the assumption of a particle possessing only magnetic charge, most of the searches are also sensitive to dyons.

**Search Techniques:** Search strategies are determined by the expected interactions of monopoles as they pass through matter. These would give rise to a number of striking characteristic signatures. Since a complete description of monopole search techniques falls outside of the scope of this minireview, only the most common methods are described below. More comprehensive descriptions of search techniques can be found in Refs. [26,27].

The induction method exploits the long-ranged electromagnetic interaction of the monopole with the quantum state of a superconducting ring which would lead to a monopole which passes through such a ring inducing a permanent current. The induction technique typically uses Superconducting Quantum Interference Devices (SQUID) technology for detection and is employed for searches for monopoles in cosmic rays and matter. Another approach is to exploit the electromagnetic energy loss of monopoles. Monopoles with Dirac charge would typically lose energy at a rate which is several thousand times larger than that expected from particles possessing the elementary electric charge. Consequently, scintillators, gas chambers and nuclear track detectors (NTDs) have been used in cosmic ray and collider experiments. A further approach, which has

been used at colliders, is to search for particles describing a non-helical path in a uniform magnetic field.

**Searches for Monopoles Bound in Matter:** Monopoles have been sought in a range of bulk materials which it is assumed would have absorbed incident cosmic ray monopoles over a long exposure time of order million years. Materials which have been studied include moon rock, meteorites, manganese modules, and sea water [28]. A stringent upper limit on the monopoles per nucleon ratio of  $\sim 10^{-29}$  has been obtained [28].

**Searches in Cosmic Rays:** Direct searches for monopoles in cosmic rays refer to those experiments in which the passage of the monopole is measured by an active detector. Catalysis processes in which GUT monopoles could induce nucleon decay are discussed in the next section. To interpret the results of the non-catalysis searches, the cross section for the catalysis process is typically either set to zero [29] or assigned a modest value (1mb) [30]. Searches which explicitly exploit the expected catalysed decays are discussed in the next section.

Although early cosmic ray searches using the induction technique [31] and NTDs [32] observed monopole candidates, none of these apparent observations have been confirmed. Recent experiments have typically employed large scale detectors. The MACRO experiment at the Gran Sasso underground laboratory comprised three different types of detector: liquid scintillator, limited stream tubes, and NTDs, which provided a total acceptance of  $\sim 10000\text{m}^2$  for an isotropic flux. As shown in Fig. 1, this experiment has so far provided the most extensive  $\beta$ -dependent flux limits for GUT monopoles with Dirac charge [30]. Also shown are limits from an experiment at the OHYA mine in Japan [29], which used a  $2000\text{m}^2$  array of NTDs.

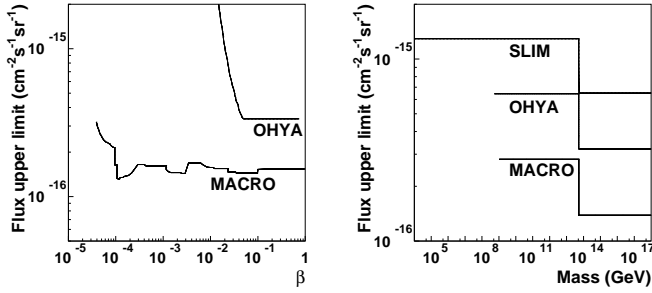
In Fig. 1, upper flux limits are also shown as a function of mass for monopole speed  $\beta > 0.05$ . In addition to MACRO and OYHA flux limits, results from the SLIM [33] high-altitude experiment are shown. The SLIM experiment provided a good sensitivity to intermediate mass monopoles ( $10^5 \lesssim M \lesssim 10^{12}$  GeV). In addition to the results shown in Fig. 1, limits as low as  $\sim 3 \times 10^{-18} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$  and  $\sim 10^{-17} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$  were obtained for monopoles with  $\beta > 0.8$  and  $\beta > 0.625$  by the IceCube [34] and Antares [35] experiments, respectively. The most stringent constraints on the flux of ultra-relativistic monopoles have been obtained by the RICE [36] and ANITA-II experiments [37] at the South Pole which were sensitive to monopoles with  $\gamma$  values of  $10^7 \lesssim \gamma \lesssim 10^{12}$  and  $10^9 \lesssim \gamma \lesssim 10^{13}$ , respectively, and which produced flux limits as low as  $10^{-19} \text{ cm}^{-2}\text{s}^{-1}\text{sr}^{-1}$ .

**Searches via the Catalysis of Nucleon-Decay:** Searches have been performed for evidence of the catalysed decay of a nucleon by a monopole, as predicted by the Callan-Rubakov mechanism. The searches are thus sensitive to the assumed value of the catalysis decay cross section. Searches have been made with the Soudan [38] and Macro [39] experiments, using tracking detectors. Searches at IMB [40], the underwater Lake Baikal

<sup>1</sup> Where no ambiguity is likely to arise, a reference to a monopole implies a particle possessing Dirac charge.

# Searches Particle Listings

## Magnetic Monopole Searches



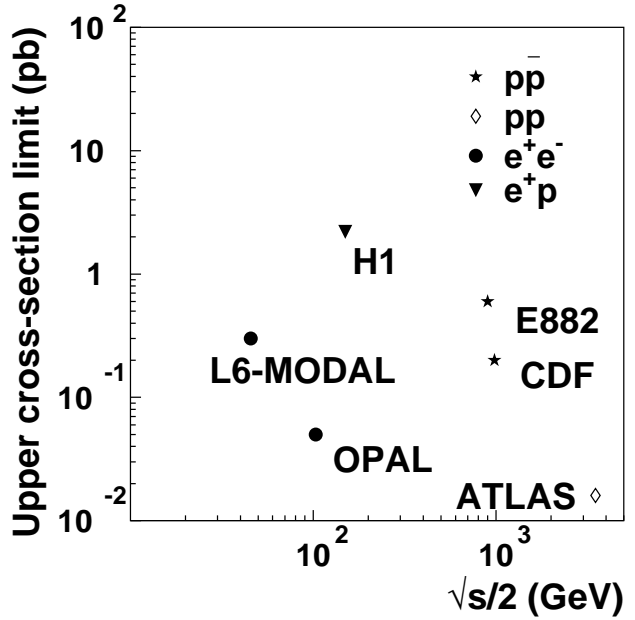
**Figure 1:** Upper flux limits for (a) GUT monopoles as a function of  $\beta$  (b) Monopoles as a function of mass for  $\beta > 0.05$ .

experiment [41] and the The IceCube experiment [42] which exploit the Cerenkov effect have also been made. The resulting  $\beta$ -dependent flux limits from these experiments typically vary between  $\sim 10^{-18}$  and  $\sim 10^{-14} \text{cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$  [25]. A recent search for low energy neutrinos (assumed to be produced from induced proton decay in the sun) was made at Super-Kamiokande [43]. A  $\beta$ -dependent limit of  $6.3 \times 10^{-24} (\frac{\beta}{10^{-3}})^2 \text{cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$  was obtained.

**Searches at Colliders:** Searches have been performed at hadron-hadron, electron-positron and lepton-hadron experiments. Collider searches can be broadly classed as being direct or indirect. In a direct search, evidence of the passage of a monopole through material, such as a charged particle track, is sought. In indirect searches, virtual monopole processes are assumed to influence the production rates of certain final states.

**Direct Searches at Colliders:** Collider experiments typically express their results in terms of upper limits on a production cross section and/or monopole mass. To calculate these limits, ansatzes are used to model the kinematics of monopole-antimonopole pair production processes since perturbative field theory cannot be used to calculate the rate and kinematic properties of produced monopoles. Limits therefore suffer from a degree of model-dependence, implying that a comparison between the results of different experiments can be problematic, in particular when this concerns excluded mass regions. A conservative approach with as little model-dependence as possible is thus to present the upper cross-section limits as a function of one half the centre-of-mass energy of the collisions, as shown in Fig. 2 for recent results from high energy colliders.

Searches for monopoles produced at the highest available energies in hadron-hadron collisions were made in  $pp$  collisions at the LHC by the ATLAS experiment [44]. In this search, highly ionising particles leaving characteristic energy deposition profiles were sought. Tevatron searches have also been carried out by the CDF [45] and E882 [46] experiments. The CDF experiment used a dedicated time-of-flight system whereas the E882 experiment employed the induction technique to search for stopped monopoles in discarded detector material which had been part of the CDF and D0 detectors using periods of luminosity. Earlier searches at the Tevatron, such as [47], used



**Figure 2:** Upper limits on the production cross sections of monopoles from various collider-based experiments.

NTDs and were based on comparatively modest amounts of integrated luminosity. Lower energy hadron-hadron experiments have employed a variety of search techniques including plastic track detectors [48] and searches for trapped monopoles [49].

The only LEP-2 search was made by OPAL [50] which quoted cross section limits for the production of monopoles possessing masses up to around 103 GeV. At LEP-1, searches were made with NTDs deployed around an interaction region. This allowed a range of charges to be sought for masses up to  $\sim 45$  GeV. The L6-MODAL experiment [51] gave limits for monopoles with charges in the range  $0.9Q_M^D$  and  $3.6Q_M^D$ , whilst an earlier search by the MODAL experiment was sensitive to monopoles with charges as low as  $0.1Q_M^D$  [52]. The deployment of NTDs around the beam interaction point was also used at earlier  $e^+e^-$  colliders such as KEK [53] and PETRA [54]. Searches at  $e^+e^-$  facilities have also been made for particles following non-helical trajectories [55,56].

There has so far been one search for monopole production in lepton-hadron scattering. Using the induction method, monopoles were sought which could have stopped in the aluminium beam pipe which had been used by the H1 experiment at HERA [57]. Cross section limits were set for monopoles with charges in the range  $Q_M^D - 6Q_M^D$  for masses up to around 140 GeV.

**Indirect Searches at Colliders:** It has been proposed that virtual monopoles can mediate processes which give rise to multi-photon final-states [58,59]. Photon-based searches were made by the D0 [60] and L3 [61] experiments. The D0 work led to spin-dependent lower mass limits of between 610 and 1580 GeV, while L3 reported a lower mass limit of 510 GeV. However,

it should be stressed that uncertainties on the theoretical calculations which were used to derive these limits are difficult to estimate.

## References

1. P.A.M.Dirac, Proc. Royal Soc. London **A133**, 60 (1931).
2. F. Englert and P. Windey, Phys. Rev. **D14**, 2728 (1976).
3. P. Goddard, J. Nuyts, and D.I. Olive, Nucl. Phys. **B125**, 1 (1977).
4. G.'t Hooft, Nucl. Phys. **B79**, 276 (1974).
5. A.M. Polyakov, Pisma Zh. Eksp. Teor. Fiz. **20**, 194 (1974) [Pisma Zh. Eksp. Teor. Fiz. **20**, 430 (1974)].
6. C.P. Dokos and T.N. Tomaras, Phys. Rev. **D21**, 2940 (1980).
7. G. Lazarides and Q. Shafi, Phys. Lett. **B94**, 149 (1980).
8. C.G. Callan, Phys. Rev. **D26**, 2058 (1982).
9. V.A. Rubakov, Nucl. Phys. **B203**, 311 (1982).
10. T.W.B. Kibble, J. Phys. **A9**, 1387 (1976).
11. J. Preskill, Phys. Rev. Lett. **43**, 1365 (1979).
12. E.N. Parker, Astrophys. J. **160**, 383 (1970).
13. M.S. Turner, E.N. Parker, and T.J. Bogdan, Phys. Rev. **D26**, 1296 (1982).
14. F.C. Adams *et al.*, Phys. Rev. Lett. **70**, 2511 (1993).
15. Y. Rephaeli and M.S. Turner, Phys. Lett. **B121**, 115 (1983).
16. E.W. Kolb, S.A. Colgate, and J.A. Harvey, Phys. Rev. Lett. **49**, 1373 (1982).
17. S. Dimopoulos, J. Preskill, and F. Wilczek, Phys. Lett. **B119**, 320 (1982).
18. K. Freese, M.S. Turner, and D.N. Schramm, Phys. Rev. Lett. **51**, 1625 (1983).
19. E.W. Kolb and M.S. Turner, Astrophys. J. **286**, 702 (1984).
20. J.A. Harvey, Nucl. Phys. **B236**, 255 (1984).
21. K. Freese and E. Krasteva, Phys. Rev. **D59**, 063007 (1999).
22. J. Arafune, M. Fukugita, and S. Yanagita, Phys. Rev. **D32**, 2586 (1985).
23. L.L. Vant-Hull, Phys. Rev. **173**, 1412 (1968).
24. S. Graf, A. Schaefer, and W. Greiner, Phys. Lett. **B262**, 463 (1991).
25. Review of Particle Physics 2012 (*this Review*), listing on *Searches for Magnetic Monopoles*.
26. G. Giacomelli and L. Patrizii, arXiv:hep-ex/0506014.
27. M. Fairbairn *et al.*, Phys. Reports **438**, 1 (2007).
28. J.M. Kovalik and J.L. Kirschvink, Phys. Rev. **A33**, 1183 (1986);  
H. Jeon and M. J. Longo, Phys. Rev. Lett. **75**, 1443 (1995) [Erratum-ibid. **76**, 159 (1996)].
29. S. Orito *et al.*, Phys. Rev. Lett. **66**, 1951 (1991).
30. M. Ambrosio *et al.* [MACRO Collab.], Eur. Phys. J. **C25**, 511 (2002).
31. B. Cabrera, Phys. Rev. Lett. **48**, 1378 (1982).
32. P.B. Price *et al.*, Phys. Rev. Lett. **35**, 487 (1975).
33. S. Balestra *et al.*, Eur. Phys. J. **C55**, 57 (2008).
34. R. Abbasi *et al.* [IceCube Collab.], Phys. Rev. **D87**, 022001 (2013).
35. S. Adrian-Martinez *et al.* [ANTARES Collab.], Astropart. Phys. **35**, 634 (2012).
36. D.P. Hogan *et al.*, Phys. Rev. **D78**, 075031 (2008).
37. M. Detrixhe *et al.*, Phys. Rev. **D83**, 023513 (2011).
38. J.E. Bartelt *et al.*, Phys. Rev. **D36**, 1990 (1987) [Erratum-ibid. **D40**, 1701 (1989)].
39. M. Ambrosio *et al.*, Eur. Phys. J. **C26**, 163 (2002).
40. R. Becker-Szendy *et al.*, Phys. Rev. **D49**, 2169 (1994).
41. V. A. Balkanov *et al.*, Prog. in Part. Nucl. Phys. **40**, 391 (1998).
42. M.G. Aartsen *et al.* [IceCube Collab.], Eur. Phys. J. **C74**, 2938 (2014).
43. K. Ueno *et al.* [Super-Kamiokande Collab.], Astropart. Phys. **36**, 131 (2012).
44. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. Lett. **109**, 261803 (2012).
45. A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **96**, 201801 (2006).
46. G.R. Kalbfleisch *et al.*, Phys. Rev. **D69**, 052002 (2004).
47. P.B. Price, G.X. Ren, and K. Kinoshita, Phys. Rev. Lett. **59**, 2523 (1987).
48. B. Aubert *et al.*, Phys. Lett. **B120**, 465 (1983).
49. R.A. Carrigan, F.A. Nezrick, and B.P. Strauss, Phys. Rev. **D8**, 3717 (1973).
50. G. Abbiendi *et al.* [OPAL Collab.], Phys. Lett. **B663**, 37 (2008).
51. J.L. Pinfold *et al.*, Phys. Lett. **B316**, 407 (1993).
52. K. Kinoshita *et al.*, Phys. Rev. **D46**, 881 (1992).
53. K. Kinoshita *et al.*, Phys. Lett. **B228**, 543 (1989).
54. P. Musset *et al.*, Phys. Lett. **B128**, 333 (1983).
55. T. Gentile *et al.* [Cleo Collab.], Phys. Rev. **D35**, 1081 (1987).
56. W. Braunschweig *et al.* [TASSO Collab.], Z. Phys. **C38**, 543 (1988).
57. A. Aktas *et al.* [H1 Collab.], Eur. Phys. J. **C41**, 133 (2005).
58. A. De Rujula, Nucl. Phys. **B435**, 257 (1995).
59. I.F. Ginzburg and A. Schiller, Phys. Rev. **D60**, 075016 (1999).
60. B. Abbott *et al.* [D0 Collab.], Phys. Rev. Lett. **81**, 524 (1998).
61. M. Acciarri *et al.* [L3 Collab.], Phys. Lett. **B345**, 609 (1995).

## Monopole Production Cross Section — Accelerator Searches

X-SECT (cm <sup>2</sup> )	MASS (GeV)	CHG (g)	ENERGY (GeV)	BEAM	DOCUMENT ID	TECN
<1.6E-38	200-1200	1	7000	$p\bar{p}$	<sup>1</sup> AAD	12CS ATLS
<5E-38	45-102	1	206	$e^+e^-$	<sup>2</sup> ABBIENDI	08 OPAL
<0.2E-36	200-700	1	1960	$p\bar{p}$	<sup>3</sup> ABULENCIA	06K CNTR
<2.E-36		1	300	$e^+p$	<sup>4,5</sup> AKTAS	05A INDU
<0.2.E-36		2	300	$e^+p$	<sup>4,5</sup> AKTAS	05A INDU
<0.09E-36		3	300	$e^+p$	<sup>4,5</sup> AKTAS	05A INDU
<0.05E-36		≥6	300	$e^+p$	<sup>4,5</sup> AKTAS	05A INDU
<2.E-36		1	300	$e^+p$	<sup>4,6</sup> AKTAS	05A INDU
<0.2E-36		2	300	$e^+p$	<sup>4,6</sup> AKTAS	05A INDU
<0.07E-36		3	300	$e^+p$	<sup>4,6</sup> AKTAS	05A INDU
<0.06E-36		≥6	300	$e^+p$	<sup>4,6</sup> AKTAS	05A INDU
<0.6E-36	>265	1	1800	$p\bar{p}$	<sup>7</sup> KALBFLEISCH	04 INDU
<0.2E-36	>355	2	1800	$p\bar{p}$	<sup>7</sup> KALBFLEISCH	04 INDU
<0.07E-36	>410	3	1800	$p\bar{p}$	<sup>7</sup> KALBFLEISCH	04 INDU
<0.2E-36	>375	6	1800	$p\bar{p}$	<sup>7</sup> KALBFLEISCH	04 INDU
<0.7E-36	>295	1	1800	$p\bar{p}$	<sup>8,9</sup> KALBFLEISCH	00 INDU
<7.8E-36	>260	2	1800	$p\bar{p}$	<sup>8,9</sup> KALBFLEISCH	00 INDU
<2.3E-36	>325	3	1800	$p\bar{p}$	<sup>8,10</sup> KALBFLEISCH	00 INDU

# Searches Particle Listings

## Magnetic Monopole Searches

< 0.11E-36	>420	6	1800	$p\overline{p}$	<sup>8,10</sup> KALBFLEISCH 00	INDU
<0.65E-33	<3.3	≥ 2	11A	<sup>197</sup> Au	<sup>11</sup> HE	97
<1.90E-33	<8.1	≥ 2	160A	<sup>208</sup> Pb	<sup>11</sup> HE	97
<3.E-37	<45.0	1.0	88-94	$e^+e^-$	PINFOLD	93 PLAS
<3.E-37	<41.6	2.0	88-94	$e^+e^-$	PINFOLD	93 PLAS
<7.E-35	<44.9	0.2-1.0	89-93	$e^+e^-$	KINOSHITA	92 PLAS
<2.E-34	<850	≥ 0.5	1800	$p\overline{p}$	BERTANI	90 PLAS
<1.2E-33	<800	≥ 1	1800	$p\overline{p}$	PRICE	90 PLAS
<1.E-37	<29	1	50-61	$e^+e^-$	KINOSHITA	89 PLAS
<1.E-37	<18	2	50-61	$e^+e^-$	KINOSHITA	89 PLAS
<1.E-38	<17	<1	35	$e^+e^-$	BRAUNSCH...	<sup>88B</sup> CNTR
<8.E-37	<24	1	50-52	$e^+e^-$	KINOSHITA	88 PLAS
<1.3E-35	<22	2	50-52	$e^+e^-$	KINOSHITA	88 PLAS
<9.E-37	<4	<0.15	10.6	$e^+e^-$	GENTILE	87 CLEO
<3.E-32	<800	≥ 1	1800	$p\overline{p}$	PRICE	87 PLAS
<3.E-38		<3	29	$e^+e^-$	FRYBERGER	84 PLAS
<1.E-31		1,3	540	$p\overline{p}$	AUBERT	<sup>83B</sup> PLAS
<4.E-38	<10	<6	34	$e^+e^-$	MUSSET	83 PLAS
<8.E-36	<20		52	$pp$	DELL	<sup>12</sup> CNTR
<9.E-37	<30	<3	29	$e^+e^-$	KINOSHITA	82 PLAS
<1.E-37	<20	<24	63	$pp$	CARRIGAN	78 CNTR
<1.E-37	<30	<3	56	$pp$	HOFFMANN	78 PLAS
			62	$pp$	DELL	76 SPRK
<4.E-33			300	$p$	<sup>12</sup> STEVENS	76B SPRK
<1.E-40	<5	<2	70	$p$	<sup>13</sup> ZRELOV	76 CNTR
<2.E-30			300	$n$	<sup>12</sup> BURKE	75 OSPK
<1.E-38			8	$\nu$	<sup>14</sup> CARRIGAN	75 HLBC
<5.E-43	<12	<10	400	$p$	EBERHARD	75B INDU
<2.E-36	<30	<3	60	$pp$	GIACOMELLI	75 PLAS
<5.E-42	<13	<24	400	$p$	CARRIGAN	74 CNTR
<6.E-42	<12	<24	300	$p$	CARRIGAN	73 CNTR
<2.E-36		1	0.001	$\gamma$	<sup>13</sup> BARTLETT	72 CNTR
<1.E-41	<5		70	$p$	GUREVICH	72 EMUL
<1.E-40	<3	<2	28	$p$	AMALDI	63 EMUL
<2.E-40	<3	<2	30	$p$	PURCELL	63 CNTR
<1.E-35	<3	<4	28	$p$	FIDECARO	61 CNTR
<2.E-35	<1	1	6	$p$	BRADNER	59 EMUL

- <sup>1</sup> AAD 12cs searched for monopoles as highly ionising objects. The cross section limits are based on an assumed Drell Yan-like production process for spin 1/2 monopoles. The limits are mass- and scenario-dependent.
- <sup>2</sup> ABBIENDI 08 assume production of spin 1/2 monopoles with effective charge  $g\beta$  ( $n=1$ ), via  $e^+e^- \rightarrow \gamma^* \rightarrow M\overline{M}$ , so that the cross section is proportional to  $(1 + \cos^2\theta)$ . There is no  $z$  information for such highly saturated tracks, so a parabolic track in the jet chamber is projected onto the  $xy$  plane. Charge per hit in the chamber produces a clean separation of signal and background.
- <sup>3</sup> ABULENCIA 06K searches for high-ionizing signals in CDF central outer tracker and time-of-flight detector. For Drell-Yan  $M\overline{M}$  production, the cross section limit implies  $M > 360$  GeV at 95% CL.
- <sup>4</sup> AKTAS 05A model-dependent limits as a function of monopole mass shown for arbitrary mass of 60 GeV. Based on search for stopped monopoles in the H1 Al beam pipe.
- <sup>5</sup> AKTAS 05A limits with assumed elastic spin 0 monopole pair production.
- <sup>6</sup> AKTAS 05A limits with assumed inelastic spin 1/2 monopole pair production.
- <sup>7</sup> KALBFLEISCH 04 reports searches for stopped magnetic monopoles in Be, Al, and Pb samples obtained from discarded material from the upgrading of DØ and CDF. A large-aperture warm-bore cryogenic detector was used. The approach was an extension of the methods of KALBFLEISCH 00. Cross section results moderately model dependent; interpretation as a mass lower limit depends on possibly invalid perturbation expansion.
- <sup>8</sup> KALBFLEISCH 00 used an induction method to search for stopped monopoles in pieces of the DØ (FNAL) beryllium beam pipe and in extensions to the drift chamber aluminum support cylinder. Results are model dependent.
- <sup>9</sup> KALBFLEISCH 00 result is for aluminum.
- <sup>10</sup> KALBFLEISCH 00 result is for beryllium.
- <sup>11</sup> HE 97 used a lead target and barium phosphate glass detectors. Cross-section limits are well below those predicted via the Drell-Yan mechanism.
- <sup>12</sup> Multiphoton events.
- <sup>13</sup> Cherenkov radiation polarization.
- <sup>14</sup> Re-examines CERN neutrino experiments.

### Monopole Production — Other Accelerator Searches

MASS (GeV)	CHG (g)	SPIN	ENERGY (GeV)	BEAM	DOCUMENT ID	TECN
> 610	≥ 1	0	1800	$p\overline{p}$	<sup>1</sup> ABBOTT 98k	D0
> 870	≥ 1	1/2	1800	$p\overline{p}$	<sup>1</sup> ABBOTT 98k	D0
>1580	≥ 1	1	1800	$p\overline{p}$	<sup>1</sup> ABBOTT 98k	D0
> 510			88-94	$e^+e^-$	<sup>2</sup> ACCIARRI 95c	L3

- <sup>1</sup> ABBOTT 98k search for heavy pointlike Dirac monopoles via central production of a pair of photons with high transverse energies.
- <sup>2</sup> ACCIARRI 95c finds a limit  $B(Z \rightarrow \gamma\gamma\gamma) < 0.8 \times 10^{-5}$  (which is possible via a monopole loop) at 95% CL and sets the mass limit via a cross section model.

### Monopole Flux — Cosmic Ray Searches

“Caty” in the charge column indicates a search for monopole-catalyzed nucleon decay.

FLUX ( $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$ )	MASS (GeV)	CHG (g)	COMMENTS ( $\beta = v/c$ )	EVTs	DOCUMENT ID	TECN
<1E-17		Caty	1E-3 < $\beta$ < 1E-2	0	<sup>1</sup> AARTSEN 14	ICCB
<3E-18		1	$\beta > 0.8$	0	<sup>2</sup> ABBASI 13	ICCB
<1.3E-17		1	$\beta > 0.625$	0	<sup>3</sup> ADRIAN-MAR..12A	ANTR
<6E-28	<1E17	Caty	1E-5 < $\beta$ < 0.04	0	<sup>4</sup> UENO 12	SKAM
<1E-19		1	$\gamma > 1E10$	0	<sup>5</sup> DETRIXHE 11	ANIT
<3.8E-17		1	$\beta > 0.76$	0	<sup>2</sup> ABBASI 10A	ICCB
<1.3E-15	1E4 < M < 5E13	1	$\beta > 0.05$	0	<sup>6</sup> BALESTRA 08	PLAS
<0.65E-15	>5E13	1	$\beta > 0.05$	0	<sup>6</sup> BALESTRA 08	PLAS
<1E-18		1	$\gamma > 1 E8$	0	<sup>5</sup> HOGAN 08	RICE
<1.4E-16		1	1.1E-4 < $\beta$ < 1	0	<sup>7</sup> AMBROSIO 02B	MCRO
<3E-16		Caty	1.1E-4 < $\beta$ < 5E-3	0	<sup>8</sup> AMBROSIO 02C	MCRO
<1.5E-15		1	5E-3 < $\beta$ < 0.99	0	<sup>9</sup> AMBROSIO 02D	MCRO
<1E-15		1	$1.1 \times 10^{-4} - 0.1$	0	<sup>10</sup> AMBROSIO 97	MCRO
<5.6E-15		1	(0.18-3.0)E-3	0	<sup>11</sup> AHLEN 94	MCRO
<2.7E-15		Caty	$\beta \sim 1 \times 10^{-3}$	0	<sup>12</sup> BECKER-SZ... 94	IMB
<8.7E-15		1	>2.E-3	0	THRON 92	SODU
<4.4E-12		1	all $\beta$	0	GARDNER 91	INDU
<7.2E-13		1	all $\beta$	0	HUBER 91	INDU
<3.7E-15	>E12	1	$\beta = 1.E-4$	0	<sup>13</sup> ORITO 91	PLAS
<3.2E-16	>E10	1	$\beta > 0.05$	0	<sup>13</sup> ORITO 91	PLAS
<3.2E-16	>E10-E12	2,3		0	<sup>13</sup> ORITO 91	PLAS
<3.8E-13		1	all $\beta$	0	BERMON 90	INDU
<5.E-16		Caty	$\beta < 1.E-3$	0	<sup>12</sup> BEZRUKOV 90	CHER
<1.8E-14		1	$\beta > 1.1E-4$	0	<sup>14</sup> BUCKLAND 90	HEPT
<1E-18		1	3.E-4 < $\beta$ < 1.5E-3	0	<sup>15</sup> GHOSH 90	MICA
<7.2E-13		1	all $\beta$	0	HUBER 90	INDU
<5.E-12	>E7	1	3.E-4 < $\beta$ < 5.E-3	0	BARISH 87	CNTR
<1.E-13		Caty	1.E-5 < $\beta$ < 1	0	<sup>12</sup> BARTELT 87	SODU
<1.E-10		1	all $\beta$	0	EBISU 87	INDU
<2.E-13		1	1.E-4 < $\beta$ < 6.E-4	0	MASEK 87	HEPT
<2.E-14		1	4.E-5 < $\beta$ < 2.E-4	0	NAKAMURA 87	PLAS
<2.E-14		1	1.E-3 < $\beta$ < 1	0	NAKAMURA 87	PLAS
<5.E-14		1	9.E-4 < $\beta$ < 1.E-2	0	SHEPKO 87	CNTR
<2.E-13		1	4.E-4 < $\beta$ < 1	0	TSUKAMOTO 87	CNTR
<5.E-14		1	all $\beta$	1	<sup>16</sup> CAPLIN 86	INDU
<5.E-12		1		0	CROMAR 86	INDU
<1.E-13		1	7.E-4 < $\beta$	0	HARA 86	CNTR
<7.E-11		1	all $\beta$	0	INCANDELA 86	INDU
<1.E-18		1	4.E-4 < $\beta$ < 1.E-3	0	<sup>15</sup> PRICE 86	MICA
<5.E-12		1		0	BERMON 85	INDU
<6.E-12		1		0	CAPLIN 85	INDU
<6.E-10		1		0	EBISU 85	INDU
<3.E-15		Caty	5.E-5 ≤ $\beta$ ≤ 1.E-3	0	<sup>12</sup> KAJITA 85	KAMI
<2.E-21		Caty	$\beta < 1.E-3$	0	<sup>12,17</sup> KAJITA 85	KAMI
<3.E-15		Caty	1.E-3 < $\beta$ < 1.E-1	0	<sup>12</sup> PARK 85B	CNTR
<5.E-12		1	1.E-4 < $\beta$ < 1	0	BATTISTONI 84	NUSX
<7.E-12		1		0	INCANDELA 84	INDU
<7.E-13		1	3.E-4 < $\beta$	0	<sup>14</sup> KAJINO 84	CNTR
<2.E-12		1	3.E-4 < $\beta$ < 1.E-1	0	KAJINO 84B	CNTR
<6.E-13		1	5.E-4 < $\beta$ < 1	0	KAWAGOE 84	CNTR
<2.E-14		1	1.E-3 < $\beta$	0	<sup>12</sup> KRISHNA... 84	CNTR
<4.E-13		1	6.E-4 < $\beta$ < 2.E-3	0	LISS 84	CNTR
<1.E-16		1	3.E-4 < $\beta$ < 1.E-3	0	<sup>15</sup> PRICE 84	MICA
<1.E-13		1	1.E-4 < $\beta$	0	PRICE 84B	PLAS
<4.E-13		1	6.E-4 < $\beta$ < 2.E-3	0	TARLE 84	CNTR
		7		0	<sup>18</sup> ANDERSON 83	EMUL
<4.E-13		1	1.E-2 < $\beta$ < 1.E-3	0	BARTELT 83B	CNTR
<1.E-12		1	7.E-3 < $\beta$ < 1	0	BARWICK 83	PLAS
<3.E-13		1	1.E-3 < $\beta$ < 4.E-1	0	BONARELLI 83	CNTR
<3.E-12		Caty	5.E-4 < $\beta$ < 5.E-2	0	<sup>12</sup> BOSETTI 83	CNTR
<4.E-11		1		0	CABRERA 83	INDU
<5.E-15		1	1.E-2 < $\beta$ < 1	0	DOKE 83	PLAS
<8.E-15		Caty	1.E-4 < $\beta$ < 1.E-1	0	<sup>12</sup> ERREDE 83	IMB
<5.E-12		1	1.E-4 < $\beta$ < 3.E-2	0	GROOM 83	CNTR
<2.E-12		1	6.E-4 < $\beta$ < 1	0	MASHIMO 83	CNTR
<1.E-13		1	$\beta = 3.E-3$	0	ALEXEYEV 82	CNTR
<2.E-12		1	7.E-3 < $\beta$ < 6.E-1	0	BONARELLI 82	CNTR
6.E-10		1	all $\beta$	1	<sup>19</sup> CABRERA 82	INDU
<2.E-11		1	1.E-2 < $\beta$ < 1.E-1	0	MASHIMO 82	CNTR
<2.E-15			concentrator	0	BARTLETT 81	PLAS
<1.E-13	>1		1.E-3 < $\beta$	0	KINOSHITA 81B	PLAS
<5.E-11	<E17		3.E-4 < $\beta$ < 1.E-3	0	ULLMAN 81	CNTR
<2.E-11			concentrator	0	BARTLETT 78	PLAS
1.E-1	>200	2		1	<sup>20</sup> PRICE 75	PLAS
<2.E-13		>2		0	FLEISCHER 71	PLAS
<1.E-19		>2	obsidian, mica	0	FLEISCHER 69C	PLAS
<5.E-15	<15	<3	concentrator	0	CARITHERS 66	ELEC
<2.E-11		<1-3	concentrator	0	MALKUS 51	EMUL

- <sup>1</sup> Beyond the monopole speed, the limits of AARTSEN 14 depend on the catalysis cross section ( $\sigma$ ) which corresponds to the monopole radiating  $\tilde{l}$  times the light per track length compared to the Cherenkov light from a single electrically charged, relativistic particle. The values quoted here correspond to  $\sigma = 1$  barn or  $\tilde{l} = 30$ .
- <sup>2</sup> ABBASI 13 and ABBASI 10A were based on a Cherenkov signature in an array of optical modules which were sunk in the Antarctic ice cap. Limits are speed-dependent.
- <sup>3</sup> ADRIAN-MARTINEZ 12A measurements were based on a Cherenkov signature in an underwater telescope in the Western Mediterranean Sea. Limits are speed-dependent.



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Searches Particle Listings  
Magnetic Monopole Searches

- <sup>4</sup> The limits from UENO 12 depend on the monopole speed and are also sensitive to assumed values of monopole mass and the catalysis cross section.
- <sup>5</sup> HOGAN 08 and DETRIXHE 11 limits on relativistic monopoles are based on nonobservation of radio Cherenkov signals at the South Pole. Limits are speed-dependent.
- <sup>6</sup> BALESTRA 08 exposed of nuclear track detector modules totaling 400 m<sup>2</sup> for 4 years at the Chacaltaya Laboratory (5230 m) in search for intermediate-mass monopoles with  $\beta > 0.05$ . The analysis is mainly based on three CR39 modules. For  $M > 5 \times 10^{13}$  GeV there can be upward-going monopoles as well, hence the flux limit is half that obtained for less massive monopoles. Previous experiments (e.g. MACRO and OHYA (ORITO 91)) had set limits only for  $M > 1 \times 10^{19}$  GeV.
- <sup>7</sup> AMBROSIO 02b direct search final result for  $m \geq 10^{17}$  GeV, based upon 4.2 to 9.5 years of running, depending upon the subsystem. Limit with CR39 track-etch detector extends the limit from  $\beta=4 \times 10^{-5}$  ( $3.1 \times 10^{-16}$  cm<sup>-2</sup> sr<sup>-1</sup> s<sup>-1</sup>) to  $\beta=1 \times 10^{-4}$  ( $2.1 \times 10^{-16}$  cm<sup>-2</sup> sr<sup>-1</sup> s<sup>-1</sup>). Limit curve in paper is piecewise continuous due to different detection techniques for different  $\beta$  ranges.
- <sup>8</sup> AMBROSIO 02c limit for catalysis of nucleon decay with catalysis cross section of  $\approx 1$  mb. The flux limit increases by  $\sim 3$  at the higher  $\beta$  limit, and increases to  $1 \times 10^{-14}$  cm<sup>-2</sup> sr<sup>-1</sup> s<sup>-1</sup> if the catalysis cross section is 0.01 mb. Based upon 71193 hr of data with the streamer detector, with an acceptance of 4250 m<sup>2</sup> sr.
- <sup>9</sup> AMBROSIO 02d result for "more than two years of data." Ionization search using several subsystems. Limit curve as a function of  $\beta$  not given. Included in AMBROSIO 02b.
- <sup>10</sup> AMBROSIO 97 global MACRO 90%CL is  $0.78 \times 10^{-15}$  at  $\beta=1.1 \times 10^{-4}$ , goes through a minimum at  $0.61 \times 10^{-15}$  near  $\beta=(1.1-2.7) \times 10^{-3}$ , then rises to  $0.84 \times 10^{-15}$  at  $\beta=0.1$ . The global limit in this region is below the Parker bound at  $10^{-15}$ . Less stringent limits are established for  $4 \times 10^{-5} < \beta < 1 \times 10^{-4}$ . Limits set by various triggers and different subdetectors are given in the paper. All limits assume a catalysis cross section smaller than a few mb.
- <sup>11</sup> AHLEN 94 limit for dyons extends down to  $\beta=0.9E-4$  and a limit of  $1.3E-4$  extends to  $\beta=0.8E-4$ . Also see comment by PRICE 94 and reply of BARISH 94. One loophole in the AHLEN 94 result is that in the case of monopoles catalyzing nucleon decay, relativistic particles could veto the events. See AMBROSIO 97 for additional results.
- <sup>12</sup> Catalysis of nucleon decay; sensitive to assumed catalysis cross section.
- <sup>13</sup> ORITO 91 limits are functions of velocity. Lowest limits are given here.
- <sup>14</sup> Used DKMPR mechanism and Penning effect.
- <sup>15</sup> Assumes monopole attaches fermion nucleus.
- <sup>16</sup> Limit from combining data of CAPLIN 86, BERMON 85, INCANDELA 84, and CABRERA 83. For a discussion of controversy about CAPLIN 86 observed event, see GUY 87. Also see SCHOUTEN 87.
- <sup>17</sup> Based on lack of high-energy solar neutrinos from catalysis in the sun.
- <sup>18</sup> Anomalous long-range  $\alpha$  (<sup>4</sup>He) tracks.
- <sup>19</sup> CABRERA 82 candidate event has single Dirac charge within  $\pm 5\%$ .
- <sup>20</sup> ALVAREZ 75, FLEISCHER 75, and FRIEDLANDER 75 explain as fragmenting nucleus. EBERHARD 75 and ROSS 76 discuss conflict with other experiments. HAGSTROM 77 reinterprets as antinucleus. PRICE 78 reassesses.

Monopole Flux — Astrophysics

FLUX (cm <sup>-2</sup> sr <sup>-1</sup> s <sup>-1</sup> )	MASS (GeV)	CHG (g)	COMMENTS ( $\beta = v/c$ )	DOCUMENT ID	TECN
<1.3E-20			faint white dwarf	<sup>1</sup> FREESE 99	ASTR
<1.E-16	E17	1	galactic field	<sup>2</sup> ADAMS 93	COSM
<1.E-23			Jovian planets	<sup>1</sup> ARAFUNE 85	ASTR
<1.E-16	E15		solar trapping	BRACCI 85b	ASTR
<1.E-18		1		<sup>1</sup> HARVEY 84	COSM
<3.E-23			neutron stars	KOLB 84	ASTR
<7.E-22			pulsars	<sup>1</sup> FREESE 83b	ASTR
<1.E-18	<E18	1	intergalactic field	<sup>1</sup> REPHAELI 83	COSM
<1.E-23			neutron stars	<sup>1</sup> DIMOPOUL... 82	COSM
<5.E-22			neutron stars	<sup>1</sup> KOLB 82	COSM
<5.E-15	>E21		galactic halo	SALPETER 82	COSM
<1.E-12	E19	1	$\beta=3.E-3$	<sup>3</sup> TURNER 82	COSM
<1.E-16		1	galactic field	PARKER 70	COSM

- <sup>1</sup> Catalysis of nucleon decay.
- <sup>2</sup> ADAMS 93 limit based on "survival and growth of a small galactic seed field" is  $10^{-16}$  ( $m/10^{17}$  GeV) cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>. Above  $10^{17}$  GeV, limit  $10^{-16}$  ( $10^{17}$  GeV/m) cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup> (from requirement that monopole density does not overclose the universe) is more stringent.
- <sup>3</sup> Re-evaluates PARKER 70 limit for GUT monopoles.

Monopole Density — Matter Searches

DENSITY	CHG (g)	MATERIAL	DOCUMENT ID	TECN
<9.8E-5/gram	$\geq 1$	Polar rock	BENDTZ 13	INDU
<6.9E-6/gram	>1/3	Meteorites and other	JEON 95	INDU
<2.E-7/gram	>0.6	Fe ore	<sup>1</sup> EBISU 87	INDU
<4.6E-6/gram	>0.5	deep schist	KOVALIK 86	INDU
<1.6E-6/gram	>0.5	manganese nodules	<sup>2</sup> KOVALIK 86	INDU
<1.3E-6/gram	>0.5	seawater	KOVALIK 86	INDU
>1.E+14/gram	>1/3	iron aerosols	MIKHAILOV 83	SPEC
<6.E-4/gram		air, seawater	CARRIGAN 76	CNTR
<5.E-1/gram	>0.04	11 materials	CABRERA 75	INDU
<2.E-4/gram	>0.05	moon rock	ROSS 73	INDU
<6.E-7/gram	<140	seawater	KOLM 71	CNTR
<1.E-2/gram	<120	manganese nodules	FLEISCHER 69	PLAS
<1.E-4/gram	>0	manganese	FLEISCHER 69b	PLAS
<2.E-3/gram	<1-3	magnetite, meteor	GOTO 63	EMUL
<2.E-2/gram		meteorite	PETUKHOV 63	CNTR

- <sup>1</sup> Mass  $1 \times 10^{14}-1 \times 10^{17}$  GeV.
- <sup>2</sup> KOVALIK 86 examined 498 kg of schist from two sites which exhibited clear mineralogical evidence of having been buried at least 20 km deep and held below the Curie temperature.

Monopole Density — Astrophysics

DENSITY	CHG (g)	MATERIAL	DOCUMENT ID	TECN
<1.E-9/gram	1	sun, catalysis	<sup>1</sup> ARAFUNE 83	COSM
<6.E-33/nuc1	1	moon wake	SCHATTEN 83	ELEC
<2.E-28/nuc1		earth heat	CARRIGAN 80	COSM
<2.E-4/prot		42cm absorption	BRODERICK 79	COSM
<2.E-13/m <sup>3</sup>		moon wake	SCHATTEN 70	ELEC

<sup>1</sup> Catalysis of nucleon decay.

REFERENCES FOR Magnetic Monopole Searches

AARTSEN 14	EPJ C74 2938	M.G. Aartsen et al.	(IceCube Collab.)
ABBASI 13	PR D87 022001	R. Abbasi et al.	(IceCube Collab.)
BENDTZ 13	PRL 110 121803	K. Bendtz et al.	
AAD 12CS	PRL 109 261803	G. Aad et al.	(ATLAS Collab.)
ADRIAN-MAR... 12A	ASP 35 634	S. Adrian-Martinez et al.	(ANTARES Collab.)
UENO 12	ASP 36 131	K. Ueno et al.	(Super-Kamiokande Collab.)
DETRIXHE 11	PR D83 023513	M. Detrixhe et al.	(ANITA Collab.)
ABBASI 10A	EPJ C69 361	R. Abbasi et al.	(IceCube Collab.)
ABBIENDI 08	PL B663 37	G. Abbiendi et al.	(OPAL Collab.)
BALESTRA 08	EPJ C55 57	S. Balestra et al.	(SLIM Collab.)
HOGAN 08	PR D78 075031	D.P. Hogan et al.	(KANS, NEBR, DELA)
ABULENCIA 05K	PRL 96 201801	A. Abulencia et al.	(CDF Collab.)
AKTAS 05A	EPJ C41 133	A. Aktas et al.	(H1 Collab.)
KALBFLEISCH 04	PR D69 052002	G.R. Kalbfleisch et al.	(OKLA)
AMBROSIO 02B	EPJ C25 511	M. Ambrosio et al.	(MACRO Collab.)
AMBROSIO 02C	EPJ C26 163	M. Ambrosio et al.	(MACRO Collab.)
AMBROSIO 02D	ASP 18 27	M. Ambrosio et al.	(MACRO Collab.)
KALBFLEISCH 00	PRL 85 5292	G.R. Kalbfleisch et al.	
FREESE 99	PR D59 063007	K. Freese, E. Krasteva	
ABBOTT 98K	PRL 81 524	B. Abbott et al.	(D0 Collab.)
AMBROSIO 97	PL B406 249	M. Ambrosio et al.	(MACRO Collab.)
HE 97	PRL 79 3134	Y.D. He	(UCB)
ACCIARRI 95C	PL B345 609	M. Acciarri et al.	(L3 Collab.)
JEON 95	PRL 75 1443	H. Jeon, M.J. Longo	(MICH)
Also	PRL 76 159 (erratum)	H. Jeon, M.J. Longo	
AHLEN 94	PRL 72 608	S.P. Ahlen et al.	(MACRO Collab.)
BARISH 94	PRL 73 1306	B.C. Barish, G. Giacomelli, J.T. Hong	(CIT+)
BECKER-SZ... 94	PR D49 2169	R.A. Becker-Szendy et al.	(IMB Collab.)
PRICE 94	PRL 73 1305	P.B. Price	(UCB)
ADAMS 93	PRL 70 2511	F.C. Adams et al.	(MICH, FNAL)
PINFOLD 93	PL B316 407	J.L. Pinfold et al.	(ALBE, HARV, MONT+)
KINOSHITA 92	PR D46 R881	K. Kinoshita et al.	(HARV, BGNA, REHO)
THRON 92	PR D46 A846	J.L. Thron et al.	(Soudan-2 Collab.)
GARDNER 91	PR D44 622	R.D. Gardner et al.	(STAN)
HUBER 91	PR D44 636	M.E. Huber et al.	(STAN)
ORITO 91	PRL 66 1951	S. Orto et al.	(ICEPP, WASCR, NIHO, ICRR)
BERMON 90	PRL 64 839	S. Bermon et al.	(IBM, BNL)
BERTANI 90	EPL 12 613	M. Bertani et al.	(BGNA, INFN)
BEZRUKOV 90	SJNP 52 54	L.B. Bezrukov et al.	(INRM)
Translated from	YAF 52 86		
BUCKLAND 90	PR D41 2726	K.N. Buckland et al.	(UCSD)
GHOSH 90	EPL 12 25	D.C. Ghosh, S. Chatterjee	(JADA)
HUBER 90	PRL 64 835	M.E. Huber et al.	(STAN)
PRICE 90	PRL 65 149	P.B. Price, J. Guir, K. Kinoshita	(UCB, HARV)
KINOSHITA 89	PL B228 543	K. Kinoshita et al.	(HARV, TISA, KEK+)
BRAUNSCH... 88B	ZPHY C38 543	R. Braunschweig et al.	(TASSO Collab.)
KINOSHITA 88	PRL 60 1610	K. Kinoshita et al.	(HARV, TISA, KEK+)
BARISH 87	PR D36 2641	B.C. Barish, G. Liu, C. Lane	(CIT)
BARTLETT 87	PR D36 1990	J.E. Bartlett et al.	(Soudan Collab.)
Also	PR D40 1701 (erratum)	J.E. Bartlett et al.	(Soudan Collab.)
EBISU 87	PR D36 3359	T. Ebisu, T. Watanabe	(KOBE)
Also	JP G11 883	T. Ebisu, T. Watanabe	(KOBE)
GENTILE 87	PR D35 1081	T. Gentile et al.	(CLEO Collab.)
GUY 87	NAT 325 463	J. Guy	(LOIC)
MASEK 87	PR D35 2758	G.E. Masek et al.	(UCSD)
NAKAMURA 87	PL B183 395	S. Nakamura et al.	(INUS, WASCR, NIHO)
PRICE 87	PRL 59 2523	P.B. Price, H. Guoxiao, K. Kinoshita	(UCB, HARV)
SCHOUTEN 87	JP E20 850	J.C. Schouten et al.	(LOIC)
SHEPKO 87	PR D35 2917	M.J. Shepko et al.	(TAMU)
TSUKAMOTO 87	EPL 3 39	T. Tsukamoto et al.	(ICRR)
CAPLIN 86	NAT 321 402	A.D. Caplin et al.	(LOIC)
Also	JP E20 850	J.C. Schouten et al.	(LOIC)
Also	NAT 325 463	J. Guy	(LOIC)
CROMAR 86	PRL 56 2561	M.W. Cromar, A.F. Clark, F.R. Fickett	(NBSB)
HARA 86	PRL 56 553	T. Hara et al.	(ICRR, KYOT, KEK, KOBE+)
INCANDELA 86	PR D34 2637	J. Incandela et al.	(CHIC, FNAL, MICH)
KOVALIK 86	PR A33 1183	J.M. Kovalik, J.L. Kirschvink	(CIT)
PRICE 86	PRL 56 1226	P.B. Price, M.H. Salamon	(UCB)
ARAFUNE 85	PR D32 2586	J. Arafune, M. Fukugita, S. Yanagita	(ICRR, KYOTU+)
BERMON 85	PRL 55 1850	S. Bermon et al.	(IBM)
BRACCI 85B	NP B258 726	L. Bracci, G. Fiorentini, G. Mezzorani	(PISA+)
Also	LNC 42 123	L. Bracci, G. Fiorentini	(PISA)
CAPLIN 85	NAT 317 234	A.D. Caplin et al.	(LOIC)
EBISU 85	JP G11 883	T. Ebisu, T. Watanabe	(LOIC)
KAJITA 85	PSJ 54 4055	T. Kajita et al.	(ICRR, KEK, NIIG)
PARK 85B	NP B252 261	H.S. Park et al.	(IMB Collab.)
BATTISTONI 84	PL 133B 454	G. Battistoni et al.	(NUS EX Collab.)
FRYBERGER 84	PR D29 1524	D. Fryberger et al.	(SLAC, UCB)
HARVEY 84	NP B236 295	J.A. Harvey	(PRIN)
INCANDELA 84	PRL 53 2067	J. Incandela et al.	(CHIC, FNAL, MICH)
KAJINO 84	PRL 52 1373	F. Kajino et al.	(ICRR)
KAJINO 84B	JP G10 447	F. Kajino et al.	(ICRR)
KAWAGOE 84	LNC 41 315	K. Kawagoe et al.	(TOKY)
KOLB 84	APJ 286 702	E.W. Kolb, M.S. Turner	(FNAL, CHIC)
KRISHNA... 84	PL 142B 99	M.R. Krishnaswamy et al.	(TATA, OSKC+)
LSS 84	PR D30 884	T.M. Liss, S.P. Ahlen, G. Tarle	(UCB, IND+)
PRICE 84	PRL 52 1265	P.B. Price et al.	(ROMA, UCB, IND+)
PRICE 84B	PL 140B 112	P.B. Price	(CERN)
TARLE 84	PRL 52 90	G. Tarle, S.P. Ahlen, T.M. Liss	(UCB, MICH+)
ANDERSON 83	PR D28 2308	S.N. Anderson et al.	(WASH)
ARAFUNE 83	PL 133B 380	J. Arafune, M. Fukugita	(ICRR, KYOTU)
AUBERT 83B	PL 120B 465	B. Aubert et al.	(CERN, LAPP)
BARTLETT 83B	PRL 50 655	J.E. Bartlett et al.	(MINN, ANL)
BARWICK 83	PR D28 2338	S.W. Barwick, K. Kinoshita, P.B. Price	(UCB)
BONARELLI 83	PL 126B 137	R. Bonarelli, P. Capiluppi, I. d'Antone	(BGNA)
BOSETTI 83	PL 133B 265	P.C. Bosetti et al.	(AACH3, HAWA, TOKY)
CABRERA 83	PRL 51 1933	B. Cabrera et al.	(STAN)
DOKE 83	PL 129B 370	T. Doke et al.	(WASU, RIKK, TTAM, RIKEN)

Searches Particle Listings  
Magnetic Monopole Searches, Supersymmetric Particle Searches

ERREDE	83	PRL 51 245	S.M. Errede <i>et al.</i>	(IMB Collab.)
FRESE	83B	PRL 51 1625	K. Freese, M.S. Turner, D.N. Schramm	(CHIC)
GROOM	83	PRL 50 573	D.E. Groom <i>et al.</i>	(UTAH, STAN)
MASHIMO	83	PL 128B 327	T. Mashimo <i>et al.</i>	(ICEPP)
MIKHAILOV	83	PL 130B 331	V.F. Mikhailov	(KAZA)
MUSSET	83	PL 128B 333	P. Musset, M. Price, E. Lohrmann	(CERN, HAMB)
REPHAEI	83	PL 121B 115	Y. Rephaeli, M.S. Turner	(CHIC)
SCHATTEN	83	PL D27 1525	K.H. Schatten	(NASA)
ALEXEYEV	82	LNC 35 413	E.N. Alekseev <i>et al.</i>	(INRM)
BONARELLI	82	PL 112B 100	R. Bonarelli <i>et al.</i>	(BGNA)
CABRERA	82	PRL 48 1378	B. Cabrera	(STAN)
DELL	82	NP B209 45	G.F. Dell <i>et al.</i>	(BNL, ADEL, ROMA)
DIMOPOUL...	82	PL 119B 320	S. Dimopoulos, J. Preskill, F. Wilczek	(HARV+)
KINOSHITA	82	PRL 48 77	K. Kinoshita, P.B. Price, D. Fryberger	(UCB+)
KOLB	82	PRL 49 1373	E.W. Kolb, S.A. Colgate, J.A. Harvey	(LASL, PRIN)
MASHIMO	82	JPS J1 3067	T. Mashimo, K. Kawagoe, M. Koshiha	(INUS)
SALPETER	82	PRL 49 1114	E.E. Salpeter, S.L. Shapiro, I. Wasserman	(CORN)
TURNER	82	PR D26 1296	M.S. Turner, E.N. Parker, T.J. Bogdan	(CHIC)
BARTLETT	81	PR D24 612	D.F. Bartlett <i>et al.</i>	(COLO, GES C)
KINOSHITA	81B	PR D24 1707	K. Kinoshita, P.B. Price	(UCB)
ULLMAN	81	PRL 47 289	J.D. Ullman	(LEHM, BNL)
CARRIGAN	80	NAT 288 348	R.A. Carrigan	(FNAL)
BRODERICK	79	PR D19 1046	J.J. Broderick <i>et al.</i>	(VPI)
BARTLETT	78	PR D18 2253	D.F. Bartlett, D. Soo, M.G. White	(COLO, PRIN)
CARRIGAN	78	PR D17 1754	R.A. Carrigan, B.P. Strauss, G. Giacomelli	(FNAL+)
HOFFMANN	78	LNC 23 357	H. Hoffmann <i>et al.</i>	(CERN, ROMA)
PRICE	78	PR D18 1382	P.B. Price <i>et al.</i>	(UCB, HOUS)
HAGSTROM	77	PRL 38 729	R. Hagstrom	(LBL)
CARRIGAN	76	PR D13 1823	R.A. Carrigan, F.A. Nezrick, B.P. Strauss	(FNAL)
DELL	76	LNC 15 269	G.F. Dell <i>et al.</i>	(CERN, BNL, ROMA, ADEL)
ROSS	76	LBL-4465	R.R. Ross	(LBL)
STEVENS	76B	PR D14 2207	D.M. Stevens <i>et al.</i>	(VPI, BNL)
ZRELOV	76	CZJP B26 1306	V.P. Zrelov <i>et al.</i>	(JINR)
ALVAREZ	75	LBL-4260	L.W. Alvarez	(LBL)
BURKE	75	PL 60B 113	D.L. Burke <i>et al.</i>	(MICH)
CABRERA	75	Thesis	B. Cabrera	(STAN)
CARRIGAN	75	NP B91 279	R.A. Carrigan, F.A. Nezrick	(FNAL)
Also		PR D3 56	R.A. Carrigan, F.A. Nezrick	(FNAL)
EBERHARD	75	PR D11 3099	P.H. Eberhard <i>et al.</i>	(LBL, MPIM)
EBERHARD	75B	LBL-4289	P.H. Eberhard	(LBL)
FLEISCHER	75	PRL 35 1412	R.L. Fleischer, R.N.F. Walker	(GESC, WUOL)
FRIEDLANDER	75	PRL 35 1167	M.W. Friedlander	(WUOL)
GIACOMELLI	75	NC 28A 21	G. Giacomelli <i>et al.</i>	(BGNA, CERN, SACL+)
PRICE	75	PRL 35 487	P.B. Price <i>et al.</i>	(UCB, HOUS)
CARRIGAN	74	PR D10 3867	R.A. Carrigan, F.A. Nezrick, B.P. Strauss	(FNAL)
CARRIGAN	73	PR D8 3717	R.A. Carrigan, F.A. Nezrick, B.P. Strauss	(FNAL)
ROSS	73	PR D8 698	R.R. Ross <i>et al.</i>	(LBL, SLAC)
Also		PR D4 3260	P.H. Eberhard <i>et al.</i>	(LBL, SLAC)
Also		SCI 167 701	L.W. Alvarez <i>et al.</i>	(LBL, SLAC)
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FLEISCHER	69C	PR 184 1398	R.L. Fleischer, P.B. Price, R.T. Woods	(GESC)
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Supersymmetric Particle Searches

SUPERSYMMETRY, PART I (THEORY)

Revised September 2015 by Howard E. Haber (UC Santa Cruz).

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**I.1. Introduction:** Supersymmetry (SUSY) is a generalization of the space-time symmetries of quantum field theory which transforms fermions into bosons and vice versa [1]. The existence of such a non-trivial extension of the Poincaré symmetry of ordinary quantum field theory was initially surprising, and its form is highly constrained by theoretical principles [2]. Supersymmetry also provides a framework for the unification of particle physics and gravity [3–6] at the Planck energy scale,  $M_P \approx 10^{19}$  GeV, where the gravitational interactions become comparable in magnitude to the gauge interactions. Moreover, supersymmetry can provide an explanation of the large hierarchy between the energy scale that characterizes electroweak symmetry breaking (of order 100 GeV) and the Planck scale [7–10]. The stability of this large *gauge hierarchy* with respect to radiative quantum corrections is not possible to maintain in the Standard Model without an *unnatural* fine-tuning of the parameters of the fundamental theory at the Planck scale. In contrast, in a supersymmetric extension of the Standard Model, it is possible to maintain the gauge hierarchy while providing a natural framework for elementary scalar fields.

If supersymmetry were an exact symmetry of nature, then particles and their *superpartners*, which differ in spin by half a unit, would be degenerate in mass. Since superpartners have not (yet) been observed, supersymmetry must be a broken symmetry. Nevertheless, the stability of the gauge hierarchy can still be maintained if the supersymmetry breaking is *soft* [11,12], and the corresponding supersymmetry-breaking mass parameters are no larger than a few TeV. Whether this is still plausible in light of recent supersymmetry searches at the LHC [13] will be discussed in Section I.7.

In particular, soft-supersymmetry-breaking terms of the Lagrangian involve combinations of fields with total mass dimension of three or less, with some restrictions on the dimension-three terms as elucidated in Ref. 11. The impact of the soft terms becomes negligible at energy scales much larger than the

size of the supersymmetry-breaking masses. Thus, a theory of *weak-scale supersymmetry*, where the effective scale of supersymmetry breaking is tied to the scale of electroweak symmetry breaking, provides a *natural* framework for the origin and the stability of the gauge hierarchy [7–10].

At present, there is no unambiguous experimental evidence for the breakdown of the Standard Model at or below the TeV scale. The expectations for new TeV-scale physics beyond the Standard Model are based primarily on three theoretical arguments. First, in a theory with an elementary scalar field of mass  $m$  and interaction strength  $\lambda$  (e.g., a quartic scalar self-coupling, the square of a gauge coupling or the square of a Yukawa coupling), the stability with respect to quantum corrections requires the existence of an energy cutoff roughly of order  $(16\pi^2/\lambda)^{1/2}m$ , beyond which new physics must enter [14]. A significantly larger energy cutoff would require an unnatural fine-tuning of parameters that govern the low-energy theory. Applying this argument to the Standard Model leads to an expectation of new physics at the TeV scale [10].

Second, the unification of the three Standard Model gauge couplings at a very high energy close to the Planck scale is possible if new physics beyond the Standard Model (which modifies the running of the gauge couplings above the electroweak scale) is present. The minimal supersymmetric extension of the Standard Model (MSSM), where superpartner masses lie below a few TeV, provides an example of successful gauge coupling unification [15].

Third, the existence of dark matter, which makes up approximately one quarter of the energy density of the universe, cannot be explained within the Standard Model of particle physics [16]. Remarkably, a stable weakly-interacting massive particle (WIMP) whose mass and interaction rate are governed by new physics associated with the TeV-scale can be consistent with the observed density of dark matter (this is the so-called *WIMP miracle*, which is reviewed in Ref. 17). The lightest supersymmetric particle, if stable, is a promising (although not the unique) candidate for the dark matter [18–22]. Further aspects of dark matter can be found in Ref. 23.

Another phenomenon not explained by the Standard Model is the origin of the matter–antimatter asymmetry of the universe [24,25]. Models of baryogenesis must satisfy the three Sakharov conditions [26]:  $C$  and  $CP$  violation, baryon number violation and a departure from thermal equilibrium. For example, the matter–antimatter asymmetry in the early universe can be generated at the electroweak phase transition if the transition is sufficiently first-order [27]. These conditions are not satisfied in the Standard Model, since the  $CP$  violation is too small and the phase transition is not strongly first-order [24,27]. In contrast, it is possible to satisfy these conditions in supersymmetric extensions of the Standard Model, where new sources of  $CP$ -violation exist and supersymmetric loops provide corrections to the temperature-dependent effective potential that can render the transition sufficiently first-order. The MSSM parameter space in which electroweak baryogenesis

occurs is strongly constrained by LHC data [28]. However, extended supersymmetric models provide new opportunities for successful electroweak baryogenesis [29]. Alternative mechanisms for baryogenesis in supersymmetric models, including the Affleck-Dine mechanism [30,24] (where a baryon asymmetry is generated through coherent scalar fields) and leptogenesis [31] (where the lepton asymmetry is converted into a baryon asymmetry at the electroweak phase transition) have also been considered in the literature.

**I.2. Structure of the MSSM:** The minimal supersymmetric extension of the Standard Model consists of the fields of the two-Higgs-doublet extension of the Standard Model and the corresponding superpartners [32,33]. A particle and its superpartner together form a supermultiplet. The corresponding field content of the supermultiplets of the MSSM and their gauge quantum numbers are shown in Table 1. The electric charge  $Q = T_3 + \frac{1}{2}Y$  is determined in terms of the third component of the weak isospin ( $T_3$ ) and the U(1) weak hypercharge ( $Y$ ).

**Table 1:** The fields of the MSSM and their  $SU(3) \times SU(2) \times U(1)$  quantum numbers are listed. For simplicity, only one generation of quarks and leptons is exhibited. For each lepton, quark, and Higgs super-multiplet, there is a corresponding anti-particle multiplet of charge-conjugated fermions and their associated scalar partners [34].

Field Content of the MSSM						
Super-multiplets	Super-field	Bosonic fields	Fermionic partners	SU(3)	SU(2)	U(1)
gluon/gluino	$\hat{V}_8$	$g$	$\tilde{g}$	8	1	0
gauge/	$\hat{V}$	$W^\pm, W^0$	$\tilde{W}^\pm, \tilde{W}^0$	1	3	0
gaugino	$\hat{V}'$	$B$	$\tilde{B}$	1	1	0
slepton/	$\hat{L}$	$(\tilde{\nu}_L, \tilde{e}_L^-)$	$(\nu, e^-)_L$	1	2	-1
lepton	$\hat{E}^c$	$\tilde{e}_R^+$	$e_L^c$	1	1	2
squark/	$\hat{Q}$	$(\tilde{u}_L, \tilde{d}_L)$	$(u, d)_L$	3	2	1/3
quark	$\hat{U}^c$	$\tilde{u}_R^*$	$u_L^c$	$\bar{3}$	1	-4/3
	$\hat{D}^c$	$\tilde{d}_R^*$	$d_L^c$	$\bar{3}$	1	2/3
Higgs/	$\hat{H}_d$	$(H_d^0, H_d^-)$	$(\tilde{H}_d^0, \tilde{H}_d^-)$	1	2	-1
higgsino	$\hat{H}_u$	$(H_u^+, H_u^0)$	$(\tilde{H}_u^+, \tilde{H}_u^0)$	1	2	1

The gauge supermultiplets consist of the gluons and their *gluino* fermionic superpartners and the  $SU(2) \times U(1)$  gauge bosons and their *gaugino* fermionic superpartners. The matter supermultiplets consist of three generations of left-handed quarks and leptons and their scalar superpartners (*squarks* and *sleptons*, collectively referred to as *sfermions*), and the corresponding antiparticles. The Higgs supermultiplets consist of two complex Higgs doublets, their *higgsino* fermionic superpartners, and the corresponding antiparticles. The enlarged Higgs sector of the MSSM constitutes the minimal structure needed to guarantee the cancellation of anomalies from the introduction of the higgsino superpartners. Moreover, without a second Higgs doublet, one cannot generate mass for both “up”-type and

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“down”-type quarks (and charged leptons) in a way consistent with the underlying supersymmetry [35–37].

In the most elegant treatment of supersymmetry, spacetime is extended to superspace which consists of the spacetime coordinates and new anticommuting fermionic coordinates  $\theta$  and  $\theta^\dagger$  [38]. Each supermultiplet is represented by a *superfield* that is a function of the superspace coordinates. The fields of a given supermultiplet (which are functions of the spacetime coordinates) are components of the corresponding superfield.

Vector superfields contain the gauge boson fields and their gaugino partners. Chiral superfields contain the spin-0 and spin-1/2 fields of the matter or Higgs supermultiplets. A general supersymmetric Lagrangian is determined by three functions of the chiral superfields [5]: the superpotential, the Kähler potential, and the gauge kinetic function (which can be appropriately generalized to accommodate higher derivative terms [39]). Minimal forms for the Kähler potential and gauge kinetic function, which generate canonical kinetic energy terms for all the fields, are required for *renormalizable* globally supersymmetric theories. A renormalizable superpotential, which is at most cubic in the chiral superfields, yields supersymmetric Yukawa couplings and mass terms. A combination of gauge invariance and supersymmetry produces couplings of gaugino fields to matter (or Higgs) fields and their corresponding superpartners. The (renormalizable) MSSM Lagrangian is then constructed by including all possible supersymmetric interaction terms (of dimension four or less) that satisfy  $SU(3) \times SU(2) \times U(1)$  gauge invariance and  $B-L$  conservation (where  $B$  = baryon number and  $L$  = lepton number). Finally, the most general soft-supersymmetry-breaking terms consistent with these symmetries are added [11,12,40].

Although the MSSM is the focus of much of this review, there is some motivation for considering non-minimal supersymmetric extensions of the Standard Model. For example, extra structure is needed to generate non-zero neutrino masses as discussed in Section I.8. In addition, in order to address some theoretical issues and tensions associated with the MSSM, it has been fruitful to introduce one additional singlet Higgs superfield. The resulting next-to-minimal supersymmetric extension of the Standard Model (NMSSM) [41] is considered further in Sections I.4–I.7 and I.9. Finally, one is always free to add additional fields to the Standard Model along with the corresponding superpartners. However, only certain choices for the new fields (*e.g.*, the addition of complete  $SU(5)$  multiplets) will preserve the successful gauge coupling unification. Some examples will be briefly mentioned in Section I.9.

**I.2.1. R-parity and the lightest supersymmetric particle:** As a consequence of  $B-L$  invariance, the MSSM possesses a multiplicative R-parity invariance, where  $R = (-1)^{3(B-L)+2S}$  for a particle of spin  $S$  [42]. This implies that all the particles of the Standard Model have even R-parity, whereas the corresponding superpartners have odd R-parity. The conservation of R-parity in scattering and decay processes has a critical impact on supersymmetric phenomenology. For example, any

initial state in a scattering experiment will involve ordinary (R-even) particles. Consequently, it follows that supersymmetric particles must be produced in pairs. In general, these particles are highly unstable and decay into lighter states. Moreover, R-parity invariance also implies that the lightest supersymmetric particle (LSP) is absolutely stable, and must eventually be produced at the end of a decay chain initiated by the decay of a heavy unstable supersymmetric particle.

In order to be consistent with cosmological constraints, a stable LSP is almost certainly electrically and color neutral [20]. Consequently, the LSP in an R-parity-conserving theory is weakly interacting with ordinary matter, *i.e.*, it behaves like a stable heavy neutrino and will escape collider detectors without being directly observed. Thus, the canonical signature for conventional R-parity-conserving supersymmetric theories is missing (transverse) energy, due to the escape of the LSP. Moreover, as noted in Section I.1 and reviewed in Refs. [21,22], the stability of the LSP in R-parity-conserving supersymmetry makes it a promising candidate for dark matter.

**I.2.2. The goldstino and gravitino:** In the MSSM, supersymmetry breaking is accomplished by including the most general renormalizable soft-supersymmetry-breaking terms consistent with the  $SU(3) \times SU(2) \times U(1)$  gauge symmetry and R-parity invariance. These terms parameterize our ignorance of the fundamental mechanism of supersymmetry breaking. If supersymmetry breaking occurs spontaneously, then a massless Goldstone fermion called the *goldstino* ( $\tilde{G}_{1/2}$ ) must exist. The goldstino would then be the LSP, and could play an important role in supersymmetric phenomenology [43].

However, the goldstino degrees of freedom are physical only in models of spontaneously-broken global supersymmetry. If supersymmetry is a local symmetry, then the theory must incorporate gravity; the resulting theory is called supergravity [44]. In models of spontaneously-broken supergravity, the goldstino is “absorbed” by the *gravitino* ( $\tilde{G}$ ) [often called  $\tilde{g}_{3/2}$  in the older literature], the spin-3/2 superpartner of the graviton, via the super-Higgs mechanism [45]. Consequently, the goldstino is removed from the physical spectrum and the gravitino acquires a mass (denoted by  $m_{3/2}$ ). If  $m_{3/2}$  is smaller than the mass of the lightest superpartner of the Standard Model particles, then the gravitino will be the LSP.

In processes with center-of-mass energy  $E \gg m_{3/2}$ , the goldstino–gravitino equivalence theorem [46] states that the interactions of the helicity  $\pm \frac{1}{2}$  gravitino (whose properties approximate those of the goldstino) dominate those of the helicity  $\pm \frac{3}{2}$  gravitino. The interactions of gravitinos with other light fields can be described by a low-energy effective Lagrangian that is determined by fundamental principles [47].

Theories in which supersymmetry breaking is independently generated by a multiplicity of sources will yield multiple goldstino states, collectively called *goldstini* [48]. One linear combination of the goldstini is identified with the exactly massless goldstino  $\tilde{G}_{1/2}$  of global supersymmetry, which is absorbed by the gravitino in local supersymmetry as described above. The

linear combinations of goldstini orthogonal to  $\tilde{G}_{1/2}$ , sometimes called pseudo-goldstinos in the literature, acquire radiatively generated masses. Theoretical and phenomenological implications of the pseudo-goldstinos are discussed further in Ref. 48.

**I.2.3. Hidden sectors and the structure of supersymmetry breaking:** It is very difficult (perhaps impossible) to construct a realistic model of spontaneously-broken weak-scale supersymmetry where the supersymmetry breaking arises solely as a consequence of the interactions of the particles of the MSSM. An alternative scheme posits a theory consisting of at least two distinct sectors: a *visible* sector consisting of the particles of the MSSM [40] and a so-called *hidden* sector where supersymmetry breaking is generated. It is often (but not always) assumed that particles of the hidden sector are neutral with respect to the Standard Model gauge group. The effects of the hidden sector supersymmetry breaking are then transmitted to the MSSM by some mechanism (often involving the mediation by particles that comprise an additional *messenger* sector). Two theoretical scenarios that exhibit this structure are gravity-mediated and gauge-mediated supersymmetry breaking.

Supergravity models provide a natural mechanism for transmitting the supersymmetry breaking of the hidden sector to the particle spectrum of the MSSM. In models of *gravity-mediated* supersymmetry breaking, gravity is the messenger of supersymmetry breaking [49–53]. More precisely, supersymmetry breaking is mediated by effects of gravitational strength (suppressed by inverse powers of the Planck mass). The soft-supersymmetry-breaking parameters arise as model-dependent multiples of the gravitino mass  $m_{3/2}$ . In this scenario,  $m_{3/2}$  is of order the electroweak-symmetry-breaking scale, while the gravitino couplings are roughly gravitational in strength [3,54]. However, such a gravitino typically plays no direct role in supersymmetric phenomenology at colliders (except perhaps indirectly in the case where the gravitino is the LSP [55]).

Under certain theoretical assumptions on the structure of the Kähler potential (the so-called sequestered form introduced in Ref. 56), supersymmetry breaking is due entirely to the super-conformal (super-Weyl) anomaly, which is common to all supergravity models [56]. In particular, gaugino masses are radiatively generated at one-loop, and squark and slepton squared-mass matrices are flavor-diagonal. In sequestered scenarios, sfermion squared-masses arise at two-loops, which implies that gluino and sfermion masses are of the same order or magnitude. This approach is called *anomaly-mediated* supersymmetry breaking (AMSB). Indeed, anomaly mediation is more generic than originally conceived, and provides a ubiquitous source of supersymmetry breaking [57]. However in the simplest formulation of AMSB as applied to the MSSM, the squared-masses of the sleptons are negative (known as the so-called tachyonic slepton problem). It may be possible to cure this fatal flaw in non-minimal extensions of the MSSM [58]. Alternatively, one can assert that anomaly mediation is not the sole source of supersymmetry breaking in the sfermion sectors. In non-sequestered scenarios, sfermion squared-masses can arise

at tree-level, in which case squark masses would be parametrically larger than the loop-suppressed gaugino masses [59].

In *gauge-mediated* supersymmetry breaking (GMSB), gauge forces transmit the supersymmetry breaking to the MSSM. A typical structure of such models involves a hidden sector where supersymmetry is broken, a messenger sector consisting of particles (messengers) with nontrivial  $SU(3) \times SU(2) \times U(1)$  quantum numbers, and the visible sector consisting of the fields of the MSSM [60–62]. The direct coupling of the messengers to the hidden sector generates a supersymmetry-breaking spectrum in the messenger sector. Supersymmetry breaking is then transmitted to the MSSM via the virtual exchange of the messenger fields. In models of *direct gauge mediation*, there is no separate hidden sector. In particular, the sector in which the supersymmetry breaking originates includes fields that carry nontrivial Standard Model quantum numbers, which allows for the direct transmission of supersymmetry breaking to the MSSM [63].

In models of gauge-mediated supersymmetry breaking, the gravitino is the LSP [18], as its mass can range from a few eV (in the case of low supersymmetry breaking scales) up to a few GeV (in the case of high supersymmetry breaking scales). In particular, the gravitino is a potential dark matter candidate (for a review and guide to the literature, see Ref. 22). Big bang nucleosynthesis also provides some interesting constraints on the gravitino and the properties of the next-to-lightest supersymmetric particle that decays into the gravitino LSP [64]. The couplings of the helicity  $\pm\frac{1}{2}$  components of  $\tilde{G}$  to the particles of the MSSM (which approximate those of the goldstino as previously noted in Section I.2.2) are significantly stronger than gravitational strength and amenable to experimental collider analyses.

The concept of a hidden sector is more general than supersymmetry. *Hidden valley* models [65] posit the existence of a hidden sector of new particles and interactions that are very weakly coupled to particles of the Standard Model. The impact of a hidden valley on supersymmetric phenomenology at colliders can be significant if the LSP lies in the hidden sector [66].

#### I.2.4. Supersymmetry and extra dimensions:

Approaches to supersymmetry breaking have also been developed in the context of theories in which the number of space dimensions is greater than three. In particular, a number of supersymmetry-breaking mechanisms have been proposed that are inherently extra-dimensional [67]. The size of the extra dimensions can be significantly larger than  $M_P^{-1}$ ; in some cases of order  $(\text{TeV})^{-1}$  or even larger [68,69].

For example, in one approach the fields of the MSSM live on some brane (a lower-dimensional manifold embedded in a higher-dimensional spacetime), while the sector of the theory that breaks supersymmetry lives on a second spatially-separated brane. Two examples of this approach are anomaly-mediated supersymmetry breaking [56] and gaugino-mediated supersymmetry breaking [70]. In both cases, supersymmetry breaking is transmitted through fields that live in the bulk (the

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higher-dimensional space between the two branes). This setup has some features in common with both gravity-mediated and gauge-mediated supersymmetry breaking (*e.g.*, a hidden and visible sector and messengers).

Alternatively, one can consider a higher-dimensional theory that is compactified to four spacetime dimensions. In this approach, supersymmetry is broken by boundary conditions on the compactified space that distinguish between fermions and bosons. This is the so-called Scherk-Schwarz mechanism [71]. The phenomenology of such models can be strikingly different from that of the usual MSSM [72].

**I.2.5. Split-supersymmetry:** If supersymmetry is not connected with the origin of the electroweak scale, it may still be possible that some remnant of the superparticle spectrum survives down to the TeV-scale or below. This is the idea of *split-supersymmetry* [73,74], in which scalar superpartners of the quarks and leptons are significantly heavier (perhaps by many orders of magnitude) than 1 TeV, whereas the fermionic superpartners of the gauge and Higgs bosons have masses on the order of 1 TeV or below. With the exception of a single light neutral scalar whose properties are practically indistinguishable from those of the Standard Model Higgs boson, all other Higgs bosons are also assumed to be very heavy. Among the supersymmetric particles, only the fermionic superpartners may be kinematically accessible at the LHC.

In models of split supersymmetry, the top squark masses cannot be arbitrarily heavy, as these parameters enter in the radiative corrections to the observed Higgs mass. In the MSSM, a Higgs boson mass of 125 GeV [75] implies an upper bound on the mass scale that characterizes the top squarks in the range of 10 to  $10^7$  TeV [76,77,78], depending on the value of the ratio of the two neutral Higgs field vacuum expectation values (although the range of upper bounds can be relaxed by further varying other relevant MSSM parameters [78]). In some approaches, gaugino masses are one-loop suppressed relative to the sfermion masses, corresponding to the so-called *mini-split* supersymmetry spectrum [77,79]. The higgsino mass scale may or may not be likewise suppressed depending on the details of the model [80].

The supersymmetry breaking required to produce such a split-supersymmetry spectrum would destabilize the gauge hierarchy, and thus would not yield an explanation for the scale of electroweak symmetry breaking. Nevertheless, models of split-supersymmetry can account for the dark matter (which is assumed to be the LSP gaugino or higgsino) and gauge coupling unification, thereby preserving two of the good features of weak-scale supersymmetry. Finally, as a consequence of the very large squark and slepton masses, the severity of the flavor and CP-violation problems alluded to at the beginning of Section I.6 are sufficiently reduced to be consistent with experimental observations.

**I.3. Parameters of the MSSM:** The parameters of the MSSM are conveniently described by considering separately the supersymmetry-conserving and the supersymmetry-breaking

sectors. A careful discussion of the conventions used here in defining the tree-level MSSM parameters can be found in Ref. 81. For simplicity, consider first the case of one generation of quarks, leptons, and their scalar superpartners.

#### I.3.1. The supersymmetry-conserving parameters:

The parameters of the supersymmetry-conserving sector consist of: (i) gauge couplings,  $g_s$ ,  $g$ , and  $g'$ , corresponding to the Standard Model gauge group  $SU(3) \times SU(2) \times U(1)$  respectively; (ii) a supersymmetry-conserving higgsino mass parameter  $\mu$ ; and (iii) Higgs-fermion Yukawa coupling constants,  $\lambda_u$ ,  $\lambda_d$ , and  $\lambda_e$ , corresponding to the coupling of one generation of left- and right-handed quarks and leptons, and their superpartners to the Higgs bosons and higgsinos. Because there is no right-handed neutrino (and its superpartner) in the MSSM as defined here, a Yukawa coupling  $\lambda_\nu$  is not included. The complex  $\mu$  parameter and Yukawa couplings enter via the most general renormalizable R-parity-conserving superpotential,

$$W = \lambda_d \hat{H}_d \hat{Q} \hat{D}^c - \lambda_u \hat{H}_u \hat{Q} \hat{U}^c + \lambda_e \hat{H}_d \hat{L} \hat{E}^c + \mu \hat{H}_u \hat{H}_d, \quad (1)$$

where the superfields are defined in Table 1 and the gauge group indices are suppressed. The reader is warned that in the literature,  $\mu$  is sometimes defined with the opposite sign to the one given in Eq. (1).

#### I.3.2. The supersymmetry-breaking parameters:

The supersymmetry-breaking sector contains the following sets of parameters: (i) three complex gaugino Majorana mass parameters,  $M_3$ ,  $M_2$ , and  $M_1$ , associated with the  $SU(3)$ ,  $SU(2)$ , and  $U(1)$  subgroups of the Standard Model; (ii) five diagonal sfermion squared-mass parameters,  $M_Q^2$ ,  $M_U^2$ ,  $M_D^2$ ,  $M_L^2$ , and  $M_E^2$ , corresponding to the five electroweak gauge multiplets, *i.e.*, superpartners of the left-handed fields  $(u, d)_L$ ,  $u_L^c$ ,  $d_L^c$ ,  $(\nu, e^-)_L$ , and  $e_L^c$ , where the superscript  $c$  indicates a charge-conjugated fermion field [34]; and (iii) three Higgs-squark-squark and Higgs-slepton-slepton trilinear interaction terms, with complex coefficients  $\lambda_u A_U$ ,  $\lambda_d A_D$ , and  $\lambda_e A_E$  (which define the so-called “ $A$ -parameters”). The inclusion of the factors of the Yukawa couplings in the definition of the  $A$ -parameters is conventional (originally motivated by a simple class of gravity-mediated supersymmetry-breaking models [3,6]). Thus, if the  $A$ -parameters as defined above are parametrically of the same order (or smaller) relative to other supersymmetry-breaking mass parameters, then only the third generation  $A$ -parameters are phenomenologically relevant. The reader is warned that the convention for the overall sign of the  $A$ -parameters varies in the literature.

Finally, we have (iv) three scalar squared-mass parameters: two of which ( $m_1^2$  and  $m_2^2$ ) are real parameters that contribute to the diagonal Higgs squared-masses, given by  $m_1^2 + |\mu|^2$  and  $m_2^2 + |\mu|^2$ , and a third that contributes to the off-diagonal Higgs squared-mass term,  $m_{12}^2 \equiv \mu B$  (which defines the complex “ $B$ -parameter”). The breaking of the electroweak symmetry  $SU(2) \times U(1)$  to  $U(1)_{EM}$  is only possible after introducing the supersymmetry-breaking Higgs squared-mass parameters.

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Minimizing the resulting tree-level Higgs scalar potential, these three squared-mass parameters can be re-expressed in terms of the two Higgs vacuum expectation values,  $\langle H_d^0 \rangle \equiv v_d/\sqrt{2}$  and  $\langle H_u^0 \rangle \equiv v_u/\sqrt{2}$  (also called  $v_1$  and  $v_2$ , respectively, in the literature), and the CP-odd Higgs mass  $m_A$  [cf. Eqs. (3) and (4) below].

Note that  $v_d^2 + v_u^2 = 4m_W^2/g^2 \simeq (246 \text{ GeV})^2$  is fixed by the  $W$  mass and the  $SU(2)$  gauge coupling, whereas the ratio

$$\tan \beta = v_u/v_d \quad (2)$$

is a free parameter. It is convenient to choose the phases of the Higgs fields such that  $m_{12}^2$  is real and non-negative. In this case, we can adopt a convention where  $0 \leq \beta \leq \pi/2$ . The tree-level conditions for the scalar potential minimum relate the diagonal and off-diagonal Higgs squared-masses in terms of  $m_Z^2 = \frac{1}{4}(g^2 + g'^2)(v_d^2 + v_u^2)$ , the angle  $\beta$  and the CP-odd Higgs mass  $m_A$ :

$$\sin 2\beta = \frac{2m_{12}^2}{m_1^2 + m_2^2 + 2|\mu|^2} = \frac{2m_{12}^2}{m_A^2}, \quad (3)$$

$$\frac{1}{2}m_Z^2 = -|\mu|^2 + \frac{m_1^2 - m_2^2 \tan^2 \beta}{\tan^2 \beta - 1}. \quad (4)$$

One must also guard against the existence of charge and/or color breaking global minima due to non-zero vacuum expectation values for the squark and charged slepton fields. This possibility can be avoided if the  $A$ -parameters are not unduly large [50,82].

Note that supersymmetry-breaking mass terms for the fermionic superpartners of scalar fields and non-holomorphic trilinear scalar interactions (*i.e.*, interactions that mix scalar fields and their complex conjugates) have not been included above in the soft-supersymmetry-breaking sector. These terms can potentially destabilize the gauge hierarchy [11] in models with gauge-singlet superfields. The latter are not present in the MSSM; hence as noted in Ref. 12, these so-called non-standard soft-supersymmetry-breaking terms are benign. However, the coefficients of these terms (which have dimensions of mass) are expected to be significantly suppressed compared to the TeV-scale in a fundamental theory of supersymmetry-breaking [83]. Consequently, we follow the usual approach and omit these terms from further consideration.

**I.3.3. MSSM-124:** The total number of independent physical parameters that define the MSSM (in its most general form) is quite large, primarily due to the soft-supersymmetry-breaking sector. In particular, in the case of three generations of quarks, leptons, and their superpartners,  $M_Q^2$ ,  $M_U^2$ ,  $M_D^2$ ,  $M_L^2$ , and  $M_E^2$  are hermitian  $3 \times 3$  matrices, and  $A_U$ ,  $A_D$ , and  $A_E$  are complex  $3 \times 3$  matrices. In addition,  $M_1$ ,  $M_2$ ,  $M_3$ ,  $B$ , and  $\mu$  are in general complex parameters. Finally, as in the Standard Model, the Higgs-fermion Yukawa couplings,  $\lambda_f$  ( $f = u, d$ , and  $e$ ), are complex  $3 \times 3$  matrices that are related to the quark and lepton mass matrices via:  $M_f = \lambda_f v_f/\sqrt{2}$ , where  $v_e \equiv v_d$  [with  $v_u$  and  $v_d$  as defined above Eq. (2)].

However, not all these parameters are physical. Some of the MSSM parameters can be eliminated by expressing interaction eigenstates in terms of the mass eigenstates, with an appropriate redefinition of the MSSM fields to remove unphysical degrees of freedom. The analysis of Ref. 84 shows that the MSSM possesses 124 independent parameters. Of these, 18 correspond to Standard Model parameters (including the QCD vacuum angle  $\theta_{\text{QCD}}$ ), one corresponds to a Higgs sector parameter (the analogue of the Standard Model Higgs mass), and 105 are genuinely new parameters of the model. The latter include: five real parameters and three  $CP$ -violating phases in the gaugino/higgsino sector, 21 squark and slepton (sfermion) masses, 36 real mixing angles to define the sfermion mass eigenstates, and 40  $CP$ -violating phases that can appear in sfermion interactions. The most general R-parity-conserving minimal supersymmetric extension of the Standard Model (without additional theoretical assumptions) will be denoted henceforth as MSSM-124 [85].

**I.4. The supersymmetric-particle spectrum:** The supersymmetric particles (*sparticles*) differ in spin by half a unit from their Standard Model partners. The superpartners of the gauge and Higgs bosons are fermions, whose names are obtained by appending “ino” to the end of the corresponding Standard Model particle name. The gluino is the color-octet Majorana fermion partner of the gluon with mass  $M_{\tilde{g}} = |M_3|$ . The superpartners of the electroweak gauge and Higgs bosons (the gauginos and higgsinos) can mix due to  $SU(2) \times U(1)$  breaking effects. As a result, the physical states of definite mass are model-dependent linear combinations of the charged and neutral gauginos and higgsinos, called *charginos* and *neutralinos*, respectively (sometimes collectively called *electroweakinos*). The neutralinos are Majorana fermions, which can generate some distinctive phenomenological signatures [86,87]. The superpartners of the quarks and leptons are spin-zero bosons: the *squarks*, charged *sleptons*, and *sneutrinos*, respectively. A complete set of Feynman rules for the sparticles of the MSSM can be found in Ref. 88. The MSSM Feynman rules also are implicitly contained in a number of Feynman diagram and amplitude generation software packages (see *e.g.*, Refs. [89–91]).

It should be noted that all mass formulae quoted below in this section are tree-level results. Radiative loop corrections will modify these results and must be included in any precision study of supersymmetric phenomenology [92]. Beyond tree level, the definition of the supersymmetric parameters becomes convention-dependent. For example, one can define physical couplings or running couplings, which differ beyond the tree level. This provides a challenge to any effort that attempts to extract supersymmetric parameters from data. The Supersymmetry Les Houches Accord (SLHA) [93] has been adopted, which establishes a set of conventions for specifying generic file structures for supersymmetric model specifications and input parameters, supersymmetric mass and coupling spectra, and

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decay tables. These provide a universal interface between spectrum calculation programs, decay packages, and high energy physics event generators.

**I.4.1. The charginos and neutralinos:** The mixing of the charged gauginos ( $\widetilde{W}^\pm$ ) and charged higgsinos ( $H_u^+$  and  $H_d^-$ ) is described (at tree-level) by a  $2 \times 2$  complex mass matrix [94–96]:

$$M_C \equiv \begin{pmatrix} M_2 & \frac{1}{\sqrt{2}} g v_u \\ \frac{1}{\sqrt{2}} g v_d & \mu \end{pmatrix}. \quad (5)$$

To determine the physical chargino states and their masses, one must perform a singular value decomposition [97,98] of the complex matrix  $M_C$ :

$$U^* M_C V^{-1} = \text{diag}(M_{\widetilde{\chi}_1^\pm}, M_{\widetilde{\chi}_2^\pm}), \quad (6)$$

where  $U$  and  $V$  are unitary matrices, and the right-hand side of Eq. (6) is the diagonal matrix of (non-negative) chargino masses. The physical chargino states are denoted by  $\widetilde{\chi}_1^\pm$  and  $\widetilde{\chi}_2^\pm$ . These are linear combinations of the charged gaugino and higgsino states determined by the matrix elements of  $U$  and  $V$  [94–96]. The chargino masses correspond to the *singular values* [97] of  $M_C$ , i.e., the positive square roots of the eigenvalues of  $M_C^\dagger M_C$ :

$$M_{\widetilde{\chi}_1^\pm, \widetilde{\chi}_2^\pm}^2 = \frac{1}{2} \{ |\mu|^2 + |M_2|^2 + 2m_W^2 \mp \sqrt{(|\mu|^2 + |M_2|^2 + 2m_W^2)^2 - 4|\mu M_2 - m_W^2 \sin 2\beta|^2} \}, \quad (7)$$

where the states are ordered such that  $M_{\widetilde{\chi}_1^\pm} \leq M_{\widetilde{\chi}_2^\pm}$ . The relative phase of  $\mu$  and  $M_2$  is physical and potentially observable.

The mixing of the neutral gauginos ( $\widetilde{B}$  and  $\widetilde{W}^0$ ) and neutral higgsinos ( $\widetilde{H}_d^0$  and  $\widetilde{H}_u^0$ ) is described (at tree-level) by a  $4 \times 4$  complex symmetric mass matrix [94,95,99,100]:

$$M_N \equiv \begin{pmatrix} M_1 & 0 & -\frac{1}{2}g'v_d & \frac{1}{2}g'v_u \\ 0 & M_2 & \frac{1}{2}g'v_d & -\frac{1}{2}g'v_u \\ -\frac{1}{2}g'v_d & \frac{1}{2}g'v_d & 0 & -\mu \\ \frac{1}{2}g'v_u & -\frac{1}{2}g'v_u & -\mu & 0 \end{pmatrix}. \quad (8)$$

To determine the physical neutralino states and their masses, one must perform a Takagi-diagonalization [97,98,101,102] of the complex symmetric matrix  $M_N$ :

$$W^T M_N W = \text{diag}(M_{\widetilde{\chi}_1^0}, M_{\widetilde{\chi}_2^0}, M_{\widetilde{\chi}_3^0}, M_{\widetilde{\chi}_4^0}), \quad (9)$$

where  $W$  is a unitary matrix and the right-hand side of Eq. (9) is the diagonal matrix of (non-negative) neutralino masses. The physical neutralino states are denoted by  $\widetilde{\chi}_i^0$  ( $i = 1, \dots, 4$ ), where the states are ordered such that  $M_{\widetilde{\chi}_1^0} \leq M_{\widetilde{\chi}_2^0} \leq M_{\widetilde{\chi}_3^0} \leq M_{\widetilde{\chi}_4^0}$ . The  $\widetilde{\chi}_i^0$  are the linear combinations of the neutral gaugino and higgsino states determined by the matrix elements of  $W$  (which is denoted by  $N^{-1}$  in Ref. 94). The neutralino masses correspond to the singular values of  $M_N$ , i.e., the positive square roots of the eigenvalues of  $M_N^\dagger M_N$ . Exact formulae for these masses can be found in Refs. [99] and [103]. A

numerical algorithm for determining the mixing matrix  $W$  has been given in Ref. 104.

If a chargino or neutralino state approximates a particular gaugino or higgsino state, it is convenient to employ the corresponding nomenclature. Specifically, if  $|M_1|$  and  $|M_2|$  are small compared to  $m_Z$  and  $|\mu|$ , then the lightest neutralino  $\widetilde{\chi}_1^0$  would be nearly a pure *photino*,  $\widetilde{\gamma}$ , the superpartner of the photon. If  $|M_1|$  and  $m_Z$  are small compared to  $|M_2|$  and  $|\mu|$ , then the lightest neutralino would be nearly a pure *bino*,  $\widetilde{B}$ , the superpartner of the weak hypercharge gauge boson. If  $|M_2|$  and  $m_Z$  are small compared to  $|M_1|$  and  $|\mu|$ , then the lightest chargino pair and neutralino would constitute a triplet of roughly mass-degenerate pure *winos*,  $\widetilde{W}^\pm$ , and  $\widetilde{W}_3^0$ , the superpartners of the weak SU(2) gauge bosons. Finally, if  $|\mu|$  and  $m_Z$  are small compared to  $|M_1|$  and  $|M_2|$ , then the lightest chargino pair and neutralino would be nearly pure *higgsino* states, the superpartners of the Higgs bosons. Each of the above cases leads to a strikingly different phenomenology.

In the NMSSM, an additional Higgs singlet superfield is added to the MSSM. This superfield comprises two real Higgs scalar degrees of freedom and an associated neutral higgsino degree of freedom. Consequently, there are five neutralino mass eigenstates that are obtained by a Takagi-diagonalization of the  $5 \times 5$  neutralino mass matrix. In many cases, the fifth neutralino state is dominated by its SU(2)  $\times$  U(1) singlet component, and thus is very weakly coupled to the Standard Model particles and their superpartners.

**I.4.2. The squarks, sleptons and sneutrinos:** For a given fermion  $f$ , there are two superpartners,  $\widetilde{f}_L$  and  $\widetilde{f}_R$ , where the  $L$  and  $R$  subscripts simply identify the scalar partners that are related by supersymmetry to the left-handed and right-handed fermions,  $f_{L,R} \equiv \frac{1}{2}(1 \mp \gamma_5)f$ , respectively. (There is no  $\widetilde{\nu}_R$  in the MSSM.) However, in general  $\widetilde{f}_L$ – $\widetilde{f}_R$  mixing is possible, in which case  $\widetilde{f}_L$  and  $\widetilde{f}_R$  are not mass eigenstates. For three generations of squarks, one must diagonalize  $6 \times 6$  matrices corresponding to the basis  $(\widetilde{q}_{iL}, \widetilde{q}_{iR})$ , where  $i = 1, 2, 3$  are the generation labels. For simplicity, only the one-generation case is illustrated in detail below. (The effects of second and third generation squark mixing can be significant and is treated in Ref. 105.)

Using the notation of the third family, the one-generation tree-level squark squared-mass matrix is given by [106]

$$\mathcal{M}^2 = \begin{pmatrix} M_Q^2 + m_q^2 + L_q & m_q X_q^* \\ m_q X_q & M_{\widetilde{R}}^2 + m_q^2 + R_q \end{pmatrix}, \quad (10)$$

where

$$X_q \equiv A_q - \mu^* (\cot \beta)^{2T_{3q}}, \quad (11)$$

and  $T_{3q} = \frac{1}{2} [-\frac{1}{2}]$  for  $q = t [b]$ . The diagonal squared-masses are governed by soft-supersymmetry-breaking squared-masses  $M_Q^2$  and  $M_{\widetilde{R}}^2 \equiv M_U^2 [M_D^2]$  for  $q = t [b]$ , the corresponding quark masses  $m_t [m_b]$ , and electroweak correction terms:

$$L_q \equiv (T_{3q} - e_q \sin^2 \theta_W) m_Z^2 \cos 2\beta, \quad R_q \equiv e_q \sin^2 \theta_W m_Z^2 \cos 2\beta, \quad (12)$$



where  $e_q = \frac{2}{3} [-\frac{1}{3}]$  for  $q = t [b]$ . The off-diagonal squark squared-masses are proportional to the corresponding quark masses and depend on  $\tan\beta$ , the soft-supersymmetry-breaking  $A$ -parameters and the higgsino mass parameter  $\mu$ . Assuming that the  $A$ -parameters are parametrically of the same order (or smaller) relative to other supersymmetry-breaking mass parameters, it then follows that  $\tilde{q}_L$ - $\tilde{q}_R$  mixing effects are small, with the possible exception of the third generation, where mixing can be enhanced by factors of  $m_t$  and  $m_b \tan\beta$ .

In the case of third generation  $\tilde{q}_L$ - $\tilde{q}_R$  mixing, the mass eigenstates (usually denoted by  $\tilde{q}_1$  and  $\tilde{q}_2$ , with  $m_{\tilde{q}_1} < m_{\tilde{q}_2}$ ) are determined by diagonalizing the  $2 \times 2$  matrix  $\mathcal{M}^2$  given by Eq. (10). The corresponding squared-masses and mixing angle are given by [106]:

$$m_{\tilde{q}_{1,2}}^2 = \frac{1}{2} \left[ \text{Tr} \mathcal{M}^2 \mp \sqrt{(\text{Tr} \mathcal{M}^2)^2 - 4 \det \mathcal{M}^2} \right],$$

$$\sin 2\theta_{\tilde{q}} = \frac{2m_q |X_q|}{m_{\tilde{q}_2}^2 - m_{\tilde{q}_1}^2}. \quad (13)$$

The one-generation results above also apply to the charged sleptons, with the obvious substitutions:  $q \rightarrow \ell$  with  $T_{3\ell} = -\frac{1}{2}$  and  $e_\ell = -1$ , and the replacement of the supersymmetry-breaking parameters:  $M_Q^2 \rightarrow M_L^2$ ,  $M_D^2 \rightarrow M_E^2$ , and  $A_q \rightarrow A_\tau$ . For the neutral sleptons,  $\tilde{\nu}_R$  does not exist in the MSSM, so  $\tilde{\nu}_L$  is a mass eigenstate.

In the case of three generations, the supersymmetry-breaking scalar-squared masses [ $M_Q^2$ ,  $M_U^2$ ,  $M_D^2$ ,  $M_L^2$ , and  $M_E^2$ ] and the  $A$ -parameters [ $A_U$ ,  $A_D$ , and  $A_E$ ] are now  $3 \times 3$  matrices as noted in Section I.3.3. The diagonalization of the  $6 \times 6$  squark mass matrices yields  $\tilde{f}_{iL}$ - $\tilde{f}_{jR}$  mixing (for  $i \neq j$ ). In practice, since the  $\tilde{f}_L$ - $\tilde{f}_R$  mixing is appreciable only for the third generation, this additional complication can often be neglected (although see Ref. 105 for examples in which the mixing between the second and third generation squarks is relevant).

**I.5. The supersymmetric Higgs sector:** Consider first the MSSM Higgs sector [36,37,107]. Despite the large number of potential  $CP$ -violating phases among the MSSM-124 parameters, the tree-level MSSM Higgs sector is automatically  $CP$ -conserving. This follows from the fact that the only potentially complex parameter ( $m_{12}^2$ ) of the MSSM Higgs potential can be chosen real and positive by rephasing the Higgs fields, in which case  $\tan\beta$  is a real positive parameter. Consequently, the physical neutral Higgs scalars are  $CP$ -eigenstates. The MSSM Higgs sector contains five physical spin-zero particles: a charged Higgs boson pair ( $H^\pm$ ), two  $CP$ -even neutral Higgs bosons (denoted by  $h^0$  and  $H^0$  where  $m_h < m_H$ ), and one  $CP$ -odd neutral Higgs boson ( $A^0$ ). The discovery of a Standard Model-like Higgs boson at the LHC with a mass of 125 GeV [75] strongly suggests that this state should be identified with  $h^0$ , although the possibility that the 125 GeV state should be identified with  $H^0$  cannot be completely ruled out [108].

In the NMSSM [41], the scalar component of the singlet Higgs superfield adds two additional neutral states to the Higgs

sector. In this model, the tree-level Higgs sector can exhibit explicit  $CP$ -violation. If  $CP$  is conserved, then the two extra neutral scalar states are  $CP$ -even and  $CP$ -odd, respectively. These states can potentially mix with the neutral Higgs states of the MSSM. If scalar states exist that are dominantly singlet, then they are weakly coupled to Standard Model gauge bosons and fermions through their small mixing with the MSSM Higgs scalars. Consequently, it is possible that one (or both) of the singlet-dominated states is considerably lighter than the Higgs boson that was observed at the LHC.

**I.5.1 The Tree-level Higgs sector:** The properties of the Higgs sector are determined by the Higgs potential, which is made up of quadratic terms [whose squared-mass coefficients were specified above Eq. (2)] and quartic interaction terms governed by dimensionless couplings. The quartic interaction terms are manifestly supersymmetric at tree level (although these are modified by supersymmetry-breaking effects at the loop level). In general, the quartic couplings arise from two sources: (i) the supersymmetric generalization of the scalar potential (the so-called “ $F$ -terms”), and (ii) interaction terms related by supersymmetry to the coupling of the scalar fields and the gauge fields, whose coefficients are proportional to the corresponding gauge couplings (the so-called “ $D$ -terms”).

In the MSSM,  $F$ -term contributions to the quartic Higgs self-couplings are absent. As a result, the strengths of the MSSM quartic Higgs interactions are fixed in terms of the gauge couplings. Due to the resulting constraint on the form of the two-Higgs-doublet scalar potential, all the tree-level MSSM Higgs-sector parameters depend only on two quantities:  $\tan\beta$  [defined in Eq. (2)] and one Higgs mass usually taken to be  $m_A$ . From these two quantities, one can predict the values of the remaining Higgs boson masses, an angle  $\alpha$  (which measures the mixture of the original  $Y = \pm 1$  Higgs doublet states in the physical  $CP$ -even neutral scalars), and the Higgs boson self-couplings. Moreover, the tree-level mass of the lighter  $CP$ -even Higgs boson is bounded,  $m_h \leq m_Z |\cos 2\beta| \leq m_Z$  [36,37]. This bound can be substantially modified when radiative corrections are included, as discussed in Section I.5.2.

In the NMSSM, the superpotential contains a trilinear term that couples the two  $Y = \pm 1$  Higgs doublet superfields and the singlet Higgs superfield. The coefficient of this term is denoted by  $\lambda$ . Consequently, the tree-level bound for the mass of the lightest  $CP$ -even MSSM Higgs boson is modified [109],

$$m_h^2 \leq m_Z^2 \cos^2 2\beta + \frac{1}{2} \lambda^2 v^2 \sin^2 2\beta, \quad (14)$$

where  $v \equiv (v_u^2 + v_d^2)^{1/2} = 246$  GeV. If one demands that  $\lambda$  should stay finite after renormalization-group evolution up to the Planck scale, then  $\lambda$  is constrained to lie below about 0.7 at the electroweak scale. However, in light of the observed Higgs mass of 125 GeV, there is some phenomenological motivation for considering larger values of  $\lambda$  [110].

The tree-level Higgs-quark and Higgs-lepton interactions of the MSSM are governed by the Yukawa couplings that were defined by the superpotential given in Eq. (1). In particular,

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the Higgs sector of the MSSM is a Type-II two-Higgs doublet model [111], in which one Higgs doublet ( $H_d$ ) couples exclusively to the right-handed down-type quark (or lepton) fields and the second Higgs doublet ( $H_u$ ) couples exclusively to the right-handed up-type quark fields. Consequently, the diagonalization of the fermion mass matrices simultaneously diagonalizes the matrix Yukawa couplings, resulting in flavor-diagonal couplings of the neutral Higgs bosons  $h^0$ ,  $H^0$  and  $A^0$  to quark and lepton pairs.

**I.5.2 The radiatively-corrected Higgs sector:** When radiative corrections are incorporated, additional parameters of the supersymmetric model enter via virtual supersymmetric particles that can appear in loops. The impact of these corrections can be significant [112]. The qualitative behavior of these radiative corrections can be most easily seen in the large top-squark mass limit, where in addition, both the splitting of the two diagonal entries and the off-diagonal entries of the top-squark squared-mass matrix [Eq. (10)] are small in comparison to the geometric mean of the two top-squark squared-masses,  $M_S^2 \equiv M_{t_1} M_{t_2}$ . In this case (assuming  $m_A > m_Z$ ), the predicted upper bound for  $m_h$  is approximately given by

$$m_h^2 \lesssim m_Z^2 \cos^2 2\beta + \frac{3g^2 m_t^4}{8\pi^2 m_W^2} \left[ \ln \left( \frac{M_S^2}{m_t^2} \right) + \frac{X_t^2}{M_S^2} \left( 1 - \frac{X_t^2}{12M_S^2} \right) \right], \quad (15)$$

where  $X_t \equiv A_t - \mu \cot \beta$  [cf. Eq. (11)] is proportional to the off-diagonal entry of the top-squark squared-mass matrix (where for simplicity,  $A_t$  and  $\mu$  are taken to be real). The Higgs mass upper limit is saturated when  $\tan \beta$  is large (i.e.,  $\cos^2 2\beta \sim 1$ ) and  $X_t = \sqrt{6} M_S$ , which defines the so-called *maximal mixing* scenario.

A more complete treatment of the radiative corrections [113] shows that Eq. (15) somewhat overestimates the true upper bound of  $m_h$ . These more refined computations, which incorporate renormalization group improvement and the leading two-loop contributions, yield  $m_h \lesssim 135$  GeV in the large  $\tan \beta$  regime (with an accuracy of a few GeV) for  $m_t = 175$  GeV and  $M_S \lesssim 2$  TeV [113].

In addition, one-loop radiative corrections can introduce  $CP$ -violating effects in the Higgs sector, which depend on some of the  $CP$ -violating phases among the MSSM-124 parameters [114]. This phenomenon is most easily understood in a scenario where  $m_A \ll M_S$  (i.e., all five physical Higgs states are significantly lighter than the supersymmetry breaking scale). In this case, one can integrate out the heavy superpartners to obtain a low-energy effective theory with two Higgs doublets. The resulting effective two-Higgs doublet model will now contain all possible Higgs self-interaction terms (both  $CP$ -conserving and  $CP$ -violating) and Higgs-fermion interactions (beyond those of Type-II) that are consistent with electroweak gauge invariance [115].

In the NMSSM, the dominant radiative correction to Eq. (14) is the same as the one given in Eq. (15). However, in contrast to the MSSM, one does not need as large a boost from the radiative corrections to achieve a Higgs mass

of 125 GeV in certain regimes of the NMSSM parameter space (e.g.,  $\tan \beta \sim 2$  and  $\lambda \sim 0.7$ ).

**I.6. Restricting the MSSM parameter freedom:** In Sections I.4 and I.5, we surveyed the parameters that comprise the MSSM-124. However, without additional restrictions on the choice of parameters, a generic parameter set within the MSSM-124 framework is not phenomenologically viable. In particular, a generic point of the MSSM-124 parameter space exhibits: (i) no conservation of the separate lepton numbers  $L_e$ ,  $L_\mu$ , and  $L_\tau$ ; (ii) unsuppressed flavor-changing neutral currents (FCNCs); and (iii) new sources of  $CP$  violation that are inconsistent with the experimental bounds.

For example, the MSSM contains many new sources of  $CP$  violation [116]. Indeed, some combinations of the complex phases of the gaugino-mass parameters, the  $A$ -parameters, and  $\mu$  must be less than on the order of  $10^{-2}$ – $10^{-3}$  to avoid generating electric dipole moments for the neutron, electron, and atoms in conflict with observed data [117–119]. The non-observation of FCNCs [120–122] places additional strong constraints on the off-diagonal matrix elements of the squark and slepton soft-supersymmetry-breaking squared-masses and  $A$ -parameters (see Section I.3.3).

The MSSM-124 is also theoretically incomplete as it provides no explanation for the fundamental origin of the supersymmetry-breaking parameters. The successful unification of the Standard Model gauge couplings at very high energies close to the Planck scale [8,74,123,124] suggests that the high-energy structure of the theory may be considerably simpler than its low-energy realization. In a top-down approach, the dynamics that governs the more fundamental theory at high energies is used to derive the effective broken-supersymmetric theory at the TeV scale. A suitable choice for the high energy dynamics is one that yields a TeV-scale theory that satisfies all relevant phenomenological constraints.

In this Section, we examine a number of theoretical frameworks that potentially yield phenomenologically viable regions of the MSSM-124 parameter space. The resulting supersymmetric particle spectrum is then a function of a relatively small number of input parameters. This is accomplished by imposing a simple structure on the soft-supersymmetry-breaking terms at a common high-energy scale  $M_X$  (typically chosen to be the Planck scale,  $M_P$ , the grand unification scale,  $M_{GUT}$ , or the messenger scale,  $M_{\text{mess}}$ ). Using the renormalization group equations, one can then derive the low-energy MSSM parameters relevant for collider physics. The initial conditions (at the appropriate high-energy scale) for the renormalization group equations depend on the mechanism by which supersymmetry breaking is communicated to the effective low energy theory.

Examples of this scenario are provided by models of gravity-mediated, anomaly mediated and gauge-mediated supersymmetry breaking, to be discussed in more detail below. In some of these approaches, one of the diagonal Higgs squared-mass parameters is driven negative by renormalization group evolution [125]. In such models, electroweak symmetry breaking is

generated radiatively, and the resulting electroweak symmetry-breaking scale is intimately tied to the scale of low-energy supersymmetry breaking.

### I.6.1. Gaugino mass relations

One prediction that arises in many grand unified supergravity models is the unification of the (tree-level) gaugino mass parameters at some high-energy scale  $M_X = M_{\text{GUT}}$  or  $M_{\text{PL}}$ :

$$M_1(M_X) = M_2(M_X) = M_3(M_X) = m_{1/2}. \quad (16)$$

Due to renormalization group running, in the one-loop approximation the effective low-energy gaugino mass parameters (at the electroweak scale) are related:

$$M_3 = (g_s^2/g^2)M_2 \simeq 3.5M_2, \quad M_1 = (5g'^2/3g^2)M_2 \simeq 0.5M_2. \quad (17)$$

Eq. (17) can also arise more generally in gauge-mediated supersymmetry-breaking models where the gaugino masses generated at the messenger scale  $M_{\text{mess}}$  (which typically lies significantly below the unification scale where the gauge couplings unify) are proportional to the corresponding squared gauge couplings at that scale.

When Eq. (17) is satisfied, the chargino and neutralino masses and mixing angles depend only on three unknown parameters: the gluino mass,  $\mu$ , and  $\tan\beta$ . It then follows that the lightest neutralino must be heavier than 46 GeV due to the non-observation of charginos at LEP [126]. If in addition  $|\mu| \gg |M_1| \gtrsim m_Z$ , then the lightest neutralino is nearly a pure bino, an assumption often made in supersymmetric particle searches at colliders. Although Eq. (17) is often assumed in many phenomenological studies, a truly model-independent approach would take the gaugino mass parameters,  $M_i$ , to be independent parameters to be determined by experiment. Indeed, an approximately massless neutralino *cannot* be ruled out at present by a model-independent analysis [127].

It is possible that the tree-level masses for the gauginos are zero. In this case, the gaugino mass parameters arise at one-loop and do not satisfy Eq. (17). For example, the gaugino masses in AMSB models arise entirely from a model-independent contribution derived from the super-conformal anomaly [56,128]. In this case, Eq. (17) is replaced (in the one-loop approximation) by:

$$M_i \simeq \frac{b_i g_i^2}{16\pi^2} m_{3/2}, \quad (18)$$

where  $m_{3/2}$  is the gravitino mass and the  $b_i$  are the coefficients of the MSSM gauge beta-functions corresponding to the corresponding U(1), SU(2), and SU(3) gauge groups,  $(b_1, b_2, b_3) = (\frac{33}{5}, 1, -3)$ . Eq. (18) yields  $M_1 \simeq 2.8M_2$  and  $M_3 \simeq -8.3M_2$ , which implies that the lightest chargino pair and neutralino comprise a nearly mass-degenerate triplet of winos,  $\tilde{W}^\pm$ ,  $\tilde{W}^0$  (cf. Table 1), over most of the MSSM parameter space. For example, if  $|\mu| \gg m_Z$ , then Eq. (18) implies that  $M_{\tilde{\chi}_1^\pm} \simeq M_{\tilde{\chi}_1^0} \simeq M_2$  [129]. The corresponding supersymmetric phenomenology differs significantly from the standard phenomenology based on Eq. (17) [130,131].

Finally, it should be noted that the unification of gaugino masses (and scalar masses) can be accidental. In particular, the energy scale where unification takes place may not be directly related to any physical scale. One version of this phenomenon has been called *mirage unification* and can occur in certain theories of fundamental supersymmetry breaking [132].

**I.6.2. The constrained MSSM: mSUGRA, CMSSM, ...** In the *minimal* supergravity (mSUGRA) framework [3–6,49–51], a form of the Kähler potential is employed that yields minimal kinetic energy terms for the MSSM fields [53]. As a result, the soft-supersymmetry-breaking parameters at the high-energy scale  $M_X$  take a particularly simple form in which the scalar squared-masses and the  $A$ -parameters are flavor-diagonal and universal [51]:

$$\begin{aligned} M_Q^2(M_X) &= M_U^2(M_X) = M_D^2(M_X) = m_0^2 \mathbf{1}, \\ M_L^2(M_X) &= M_E^2(M_X) = m_0^2 \mathbf{1}, \\ m_1^2(M_X) &= m_2^2(M_X) = m_0^2, \\ A_U(M_X) &= A_D(M_X) = A_E(M_X) = A_0 \mathbf{1}, \end{aligned} \quad (19)$$

where  $\mathbf{1}$  is a  $3 \times 3$  identity matrix in generation space. As in the Standard Model, this approach exhibits minimal flavor violation [133,134], whose unique source is the nontrivial flavor structure of the Higgs-fermion Yukawa couplings. The gaugino masses are also unified according to Eq. (16).

Renormalization group evolution is then used to derive the values of the supersymmetric parameters at the low-energy (electroweak) scale. For example, to compute squark masses, one must use the *low-energy* values for  $M_Q^2$ ,  $M_U^2$ , and  $M_D^2$  in Eq. (10). Through the renormalization group running with boundary conditions specified in Eqs. (17) and (19), one can show that the low-energy values of  $M_Q^2$ ,  $M_U^2$ , and  $M_D^2$  depend primarily on  $m_0^2$  and  $m_{1/2}^2$ . A number of useful approximate analytic expressions for superpartner masses in terms of the mSUGRA parameters can be found in Ref. 135.

In the mSUGRA approach, four flavors of squarks (with two squark eigenstates per flavor) are nearly mass-degenerate. If  $\tan\beta$  is not very large,  $\tilde{b}_R$  is also approximately degenerate in mass with the first two generations of squarks. The  $\tilde{b}_L$  mass and the diagonal  $\tilde{t}_L$  and  $\tilde{t}_R$  masses are typically reduced relative to the common squark mass of the first two generations. In addition, there are six flavors of nearly mass-degenerate sleptons (with two slepton eigenstates per flavor for the charged sleptons and one per flavor for the sneutrinos); the sleptons are expected to be somewhat lighter than the mass-degenerate squarks. As noted below Eq. (10), third-generation squark masses and tau-slepton masses are sensitive to the strength of the respective  $\tilde{f}_L$ – $\tilde{f}_R$  mixing. The LSP is typically the lightest neutralino,  $\tilde{\chi}_1^0$ , which is dominated by its bino component. Regions of the mSUGRA parameter space in which the LSP is electrically charged do exist but are not phenomenologically viable [20].

One can count the number of independent parameters in the mSUGRA framework. In addition to 18 Standard Model

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parameters (excluding the Higgs mass), one must specify  $m_0$ ,  $m_{1/2}$ ,  $A_0$ , the Planck-scale values for  $\mu$  and  $B$ -parameters (denoted by  $\mu_0$  and  $B_0$ ), and the gravitino mass  $m_{3/2}$ . Without additional model assumptions,  $m_{3/2}$  is independent of the parameters that govern the mass spectrum of the superpartners of the Standard Model [51]. In principle,  $A_0$ ,  $B_0$ ,  $\mu_0$ , and  $m_{3/2}$  can be complex, although in the mSUGRA approach, these parameters are taken (arbitrarily) to be real.

As previously noted, renormalization group evolution is used to compute the low-energy values of the mSUGRA parameters, which then fixes all the parameters of the low-energy MSSM. In particular, the two Higgs vacuum expectation values (or equivalently,  $m_Z$  and  $\tan\beta$ ) can be expressed as a function of the Planck-scale supergravity parameters. The simplest procedure is to remove  $\mu_0$  and  $B_0$  in favor of  $m_Z$  and  $\tan\beta$  [the sign of  $\mu_0$ , denoted  $\text{sgn}(\mu_0)$  below, is not fixed in this process]. In this case, the MSSM spectrum and its interaction strengths are determined by five parameters:

$$m_0, A_0, m_{1/2}, \tan\beta, \text{ and } \text{sgn}(\mu_0), \quad (20)$$

and an independent gravitino mass  $m_{3/2}$  (in addition to the 18 parameters of the Standard Model). In Ref. 136, this framework was dubbed the *constrained minimal supersymmetric extension of the Standard Model* (CMSSM).

In the early literature, additional conditions were obtained by assuming a simplified form for the hidden sector that provides the fundamental source of supersymmetry breaking. Two additional relations emerged among the mSUGRA parameters [49,53]:  $B_0 = A_0 - m_0$  and  $m_{3/2} = m_0$ . These relations characterize a theory that was called minimal supergravity when first proposed. In the subsequent literature, it has been more common to omit these extra conditions in defining the mSUGRA model (in which case the mSUGRA model and the CMSSM are synonymous). The authors of Ref. 137 advocate restoring the original nomenclature in which the mSUGRA model is defined with the extra conditions as originally proposed. Additional mSUGRA variations can be considered where different relations among the CMSSM parameters are imposed.

One can also relax the universality of scalar masses by decoupling the squared-masses of the Higgs bosons and the squarks/sleptons. This leads to the non-universal Higgs mass models (NUHMs), thereby adding one or two new parameters to the CMSSM depending on whether the diagonal Higgs scalar squared-mass parameters ( $m_1^2$  and  $m_2^2$ ) are set equal (NUHM1) or taken to be independent (NUHM2) at the high energy scale  $M_X^2$ . Clearly, this modification preserves the minimal flavor violation of the mSUGRA approach. Nevertheless, the mSUGRA approach and its NUHM generalizations are probably too simplistic. Theoretical considerations suggest that the universality of Planck-scale soft-supersymmetry-breaking parameters is not generic [138]. In particular, effective operators at the Planck scale exist that do not respect flavor universality, and it is difficult to find a theoretical principle that would forbid them.

In the framework of supergravity, if anomaly mediation is the sole source of supersymmetry breaking, then the gaugino mass parameters, diagonal scalar squared-mass parameters, and the supersymmetry-breaking trilinear scalar interaction terms (proportional to  $\lambda_f A_F$ ) are determined in terms of the beta functions of the gauge and Yukawa couplings and the anomalous dimensions of the squark and slepton fields [56,128,131]. As noted in Section I.2.3, this approach yields tachyonic sleptons in the MSSM unless additional sources of supersymmetry breaking are present. In the *minimal* AMSB (mAMSB) scenario, a universal squared-mass parameter,  $m_0^2$ , is added to the AMSB expressions for the diagonal scalar squared-masses [131]. Thus, the mAMSB spectrum and its interaction strengths are determined by four parameters,  $m_0^2$ ,  $m_{3/2}$ ,  $\tan\beta$  and  $\text{sgn}(\mu_0)$ .

The mAMSB scenario appears to be ruled out based on the observed value of the Higgs boson mass, assuming an upper limit on  $M_S$  of a few TeV, since the mAMSB constraint on  $A_F$  implies that the maximal mixing scenario cannot be achieved [cf. Eq. (15)]. Indeed, under the stated assumptions, the mAMSB Higgs mass upper bound lies below the observed Higgs mass value [139]. Thus within the AMSB scenario, either an additional supersymmetry-breaking contribution to  $\lambda_f A_F$  and/or new ingredients beyond the MSSM are required.

**I.6.3. Gauge-mediated supersymmetry breaking:** In contrast to models of gravity-mediated supersymmetry breaking, the universality of the fundamental soft-supersymmetry-breaking squark and slepton squared-mass parameters is guaranteed in gauge-mediated supersymmetry breaking (GMSB) because the supersymmetry breaking is communicated to the sector of MSSM fields via gauge interactions [61,62]. In GMSB models, the mass scale of the messenger sector (or its equivalent) is sufficiently below the Planck scale such that the additional supersymmetry-breaking effects mediated by supergravity can be neglected.

In the minimal GMSB approach, there is one effective mass scale,  $\Lambda$ , that determines all low-energy scalar and gaugino mass parameters through loop effects, while the resulting  $A$ -parameters are suppressed. In order that the resulting superpartner masses be of order 1 TeV or less, one must have  $\Lambda \sim 100$  TeV. The origin of the  $\mu$  and  $B$ -parameters is quite model-dependent, and lies somewhat outside the ansatz of gauge-mediated supersymmetry breaking.

The simplest GMSB models appear to be ruled out based on the observed value of the Higgs boson mass. Due to suppressed  $A$  parameters, it is difficult to boost the contributions of the radiative corrections in Eq. (15) to obtain a Higgs mass as large as 125 GeV. However, this conflict can be alleviated in more complicated GMSB models [140]. To analyze these generalized GMSB models, it has been especially fruitful to develop model-independent techniques that encompass all known GMSB models [141]. These techniques are well-suited for a comprehensive analysis [142] of the phenomenological profile of gauge-mediated supersymmetry breaking.

The gravitino is the LSP in GMSB models, as noted in Section I.2.3. As a result, the next-to-lightest supersymmetric particle (NLSP) now plays a crucial role in the phenomenology of supersymmetric particle production and decays. Note that unlike the LSP, the NLSP can be charged. In GMSB models, the most likely candidates for the NLSP are  $\tilde{\chi}_1^0$  and  $\tilde{\tau}_R^\pm$ . The NLSP will decay into its superpartner plus a gravitino (*e.g.*,  $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$ ,  $\tilde{\chi}_1^0 \rightarrow Z \tilde{G}$ ,  $\tilde{\chi}_1^0 \rightarrow h^0 \tilde{G}$  or  $\tilde{\tau}_R^\pm \rightarrow \tau^\pm \tilde{G}$ ), with lifetimes and branching ratios that depend on the model parameters. There are also GMSB scenarios in which there are several nearly degenerate co-NLSP's, any one of which can be produced at the penultimate step of a supersymmetric decay chain [143]. For example, in the slepton co-NLSP case, all three right-handed sleptons are close enough in mass and thus can each play the role of the NLSP.

Different choices for the identity of the NLSP and its decay rate lead to a variety of distinctive supersymmetric phenomenologies [62,144]. For example, a long-lived  $\tilde{\chi}_1^0$ -NLSP that decays outside collider detectors leads to supersymmetric decay chains with missing energy in association with leptons and/or hadronic jets (this case is indistinguishable from the standard phenomenology of the  $\tilde{\chi}_1^0$ -LSP). On the other hand, if  $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$  is the dominant decay mode, and the decay occurs inside the detector, then nearly *all* supersymmetric particle decay chains would contain a photon. In contrast, in the case of a  $\tilde{\tau}_R^\pm$ -NLSP, the  $\tilde{\tau}_R^\pm$  would either be long-lived or would decay inside the detector into a  $\tau$ -lepton plus missing energy.

In GMSB models based on the MSSM, the fundamental origins of the  $\mu$  and  $B$ -parameters are not explicitly given, as previously noted. An alternative approach is to consider GMSB models based on the NMSSM [145]. The vacuum expectation value of the additional singlet Higgs superfield can be used to generate effective  $\mu$  and  $B$ -parameters [146]. Such models provide an alternative GMSB framework for achieving a Higgs mass of 125 GeV, while still being consistent with LHC bounds on supersymmetric particle masses [147].

**I.6.4. The phenomenological MSSM:** Of course, any of the theoretical assumptions described in this Section must be tested experimentally and could turn out to be wrong. To facilitate the exploration of MSSM phenomena in a more model-independent way while respecting the constraints noted at the beginning of this Section, the phenomenological MSSM (pMSSM) has been introduced [148].

The pMSSM is governed by 19 independent real supersymmetric parameters: the three gaugino mass parameters  $M_1$ ,  $M_2$  and  $M_3$ , the Higgs sector parameters  $m_A$  and  $\tan\beta$ , the Higgsino mass parameter  $\mu$ , five sfermion squared-mass parameters for the degenerate first and second generations ( $M_{\tilde{Q}}^2$ ,  $M_{\tilde{U}}^2$ ,  $M_{\tilde{D}}^2$ ,  $M_{\tilde{L}}^2$  and  $M_{\tilde{E}}^2$ ), the five corresponding sfermion squared-mass parameters for the third generation, and three third-generation  $A$ -parameters ( $A_t$ ,  $A_b$  and  $A_\tau$ ). As previously noted, the first and second generation  $A$ -parameters can be neglected as their phenomenological consequences are negligible.

A comprehensive study of the 19-parameter pMSSM is computationally expensive. This is somewhat ameliorated in Ref. 149, where the number of pMSSM parameters is reduced to ten by assuming one common squark squared-mass parameter for the first two generations, a second common squark squared-mass parameter for the third generation, a common slepton squared-mass parameter and a common third generation  $A$  parameter. Applications of the pMSSM approach to supersymmetric particle searches, and a discussion of the implications for past and future LHC studies can be found in Refs. [149] and [150].

#### I.6.5. Simplified models:

It is possible to focus on a small subset of the supersymmetric particle spectrum and study its phenomenology with minimal theoretical bias. In this *simplified model* approach [151], one considers the production of a pair of specific superpartners and follows their decay chains under the assumption that a limited number of decay modes dominate. Simplified models depend only on a few relevant quantities (cross sections, branching ratios and masses), and thus provide a framework for studies of supersymmetric phenomena, independently of the precise details of the theory that govern the supersymmetric parameters.

Applications of the simplified models approach to supersymmetric particle searches and a discussion of their limitations can be found in Ref. 13.

#### I.7. Experimental data confronts the MSSM:

At present, there is no evidence for weak-scale supersymmetry from the data analyzed by the LHC experiments. Recent LHC data has been especially effective in ruling out the existence of colored supersymmetric particles (primarily the gluino and the first generation of squarks) with masses below about 1 TeV [13,152]. The precise mass limits are model dependent. For example, higher mass colored superpartners have been ruled out in the context of the CMSSM. In less constrained frameworks of the MSSM, regions of parameter space can be identified in which lighter squarks and gluinos below 1 TeV cannot be definitely ruled out [13]. Additional constraints arise from limits on the contributions of virtual supersymmetric particle exchange to a variety of Standard Model processes [120–122].

In light of these negative results, one must confront the tension that exists between the theoretical expectations for the magnitude of the supersymmetry-breaking parameters and the non-observation of supersymmetric phenomena.

##### I.7.1 Naturalness constraints and the little hierarchy:

In Section I, weak-scale supersymmetry was motivated as a natural solution to the hierarchy problem, which could provide an understanding of the origin of the electroweak symmetry-breaking scale without a significant fine-tuning of the fundamental parameters that govern the MSSM. In this context, the soft-supersymmetry-breaking masses must be generally of the order of 1 TeV or below [153]. This requirement is most easily seen in the determination of  $m_Z$  by the scalar potential minimum condition. In light of Eq. (4), to avoid the fine-tuning of

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MSSM parameters, the soft-supersymmetry-breaking squared-masses  $m_1^2$  and  $m_2^2$  and the higgsino squared-mass  $|\mu|^2$  should all be roughly of  $\mathcal{O}(m_Z^2)$ . Many authors have proposed quantitative measures of fine-tuning [153–155]. One of the simplest measures is the one given by Barbieri and Giudice [153],

$$\Delta_i \equiv \left| \frac{\partial \ln m_Z^2}{\partial \ln p_i} \right|, \quad \Delta \equiv \max_i \Delta_i, \quad (21)$$

where the  $p_i$  are the MSSM parameters at the high-energy scale  $M_X$ , which are set by the fundamental supersymmetry-breaking dynamics. The theory is more fine-tuned as  $\Delta$  becomes larger.

One can apply the fine-tuning measure to any explicit model of supersymmetry breaking. For example, in the approaches discussed in Section I.6, the  $p_i$  are parameters of the model at the energy scale  $M_X$  where the soft-supersymmetry-breaking operators are generated by the dynamics of supersymmetry breaking. Renormalization group evolution then determines the values of the parameters appearing in Eq. (4) at the electroweak scale. In this way,  $\Delta$  is sensitive to all the supersymmetry-breaking parameters of the model (see e.g. Ref. 156).

As anticipated, there is a tension between the present experimental lower limits on the masses of colored supersymmetric particles [157,158] and the expectation that supersymmetry-breaking is associated with the electroweak symmetry-breaking scale. Moreover, this tension is exacerbated by the observed value of the Higgs mass ( $m_h \simeq 125$  GeV), which is not far from the MSSM upper bound ( $m_h \lesssim 135$  GeV) [which depends on the top-squark mass and mixing as noted in Section I.5.2]. If  $M_{\text{SUSY}}$  characterizes the scale of supersymmetric particle masses, then one would crudely expect  $\Delta \sim M_{\text{SUSY}}^2/m_Z^2$ . For example, if  $M_{\text{SUSY}} \sim 1$  TeV then there must be at least a  $\Delta^{-1} \sim 1\%$  fine-tuning of the MSSM parameters to achieve the observed value of  $m_Z$ . This separation of the electroweak symmetry-breaking and supersymmetry-breaking scales is an example of the *little hierarchy problem* [159,160].

However, one must be very cautious when drawing conclusions about the viability of weak-scale supersymmetry to explain the origin of electroweak symmetry breaking [161]. First, one must decide the largest tolerable value of  $\Delta$  within the framework of weak-scale supersymmetry (should it be  $\Delta \sim 10^2$ – $100^3$ ?). Second, the fine-tuning parameter  $\Delta$  depends quite sensitively on the structure of the supersymmetry-breaking dynamics, such as the value of  $M_X$  and relations among supersymmetry-breaking parameters in the fundamental high energy theory [162]. For example, in so-called focus point supersymmetry models [163], all squark masses can be as heavy as 5 TeV *without* significant fine-tuning. This can be attributed to a focusing behavior of the renormalization group evolution where certain relations hold among the high-energy values of the scalar squared-mass supersymmetry-breaking parameters.

Among the colored superpartners, the third generation squarks generically have the most significant impact on the naturalness constraints [164], while their masses are the least constrained by the LHC data. Hence, in the absence of any

relation between third generation squarks and those of the first two generations, the naturalness constraints due to present LHC data can be considerably weaker than those obtained in the CMSSM. Indeed, models with first and second generation squark masses in the multi-TeV range do not generically require significant fine tuning. Such models have the added benefit that undesirable FCNCs mediated by squark exchange are naturally suppressed [165]. Other MSSM mass spectra that are compatible with moderate fine tuning have been considered in Refs. [152,162,166].

The lower bounds on squark and gluino masses may not be as large as suggested by the experimental analyses based on the CMSSM or simplified models. For example, mass bounds for the gluino and the first and second generation squarks based on the CMSSM can often be evaded in alternative or extended MSSM models, *e.g.*, compressed supersymmetry [167] and stealth supersymmetry [168]. Moreover, experimental limits on the masses for the third generation squarks (which enter the fine-tuning considerations more directly) are less constrained than the masses of other colored supersymmetric states.

Among the uncolored superpartners, the higgsinos are the most impacted by the naturalness constraints. In light of Eq. (4), the masses of the two neutral higgsinos and charged higgsino pair (which are governed by  $|\mu|$ ) should not be significantly larger than  $m_Z$  to avoid an unnatural fine-tuning of the supersymmetric parameters. The experimental limits on the masses of such light higgsinos are not well constrained, as they are difficult to detect directly at the LHC due to their soft decay products.

Finally, one can also consider extensions of the MSSM in which the degree of fine-tuning is relaxed. For example, it has already been noted in Section I.5.2 that it is possible to accommodate the observed Higgs mass more easily in the NMSSM due to contributions to  $m_h^2$  proportional to the parameter  $\lambda$ . This means that we do not have to rely on a large contribution from the radiative corrections to boost the Higgs mass sufficiently above its tree-level bound. This allows for smaller top squark masses, which are more consistent with the demands of naturalness. The reduction of the fine-tuning in various NMSSM models was initially advocated in Ref. 169, and more recently has been exhibited in Refs. [110,170]. Naturalness can also be relaxed in extended supersymmetric models with vector-like quarks [171] and in gauge extensions of the MSSM [172].

Thus, it is premature to conclude that weak-scale supersymmetry is on the verge of exclusion. Nevertheless, it is possible to sharpen the upper bounds on superpartner masses based on naturalness arguments, which ultimately will either confirm or refute the weak scale supersymmetry hypothesis [173].

#### I.7.2 Constraints from virtual exchange of supersymmetric particles

There are a number of low-energy measurements that are sensitive to the effects of new physics through indirect searches

via supersymmetric loop effects. For example, the virtual exchange of supersymmetric particles can contribute to the muon anomalous magnetic moment,  $a_\mu \equiv \frac{1}{2}(g-2)_\mu$ , as reviewed in Ref. 174. The Standard Model prediction for  $a_\mu$  exhibits a deviation at the level of  $3-4\sigma$  from the experimentally observed value [175]. This discrepancy is difficult to accommodate in the constrained supersymmetry models of Section I.6.2 and I.6.3 given the present sparticle mass bounds [158]. Nevertheless, there are regions of the more general pMSSM parameter space that are consistent with the observed value of  $a_\mu$  [176].

The rare inclusive decay  $b \rightarrow s\gamma$  also provides a sensitive probe to the virtual effects of new physics beyond the Standard Model. Recent experimental measurements of  $B \rightarrow X_s + \gamma$  [177] are in very good agreement with the theoretical Standard Model predictions of Ref. 178. Since supersymmetric loop corrections can contribute an observable shift from the Standard Model predictions, the absence of any significant deviations places useful constraints on the MSSM parameter space [179].

The rare decay  $B_s \rightarrow \mu^+\mu^-$  is especially sensitive to supersymmetric loop effects, with some loop contributions scaling as  $\tan^6\beta$  when  $\tan\beta \gg 1$  [180]. The observation of this rare decay mode along with the first observation of  $B_d \rightarrow \mu^+\mu^-$  are compatible with the predicted Standard Model rates at the  $1.2\sigma$  and  $2.2\sigma$  level, respectively [181].

The decays  $B^\pm \rightarrow \tau^\pm\nu_\tau$  and  $B \rightarrow D^{(*)}\tau^-\bar{\nu}_\tau$  are noteworthy, since in models with extended Higgs sectors such as the MSSM, these processes possess *tree-level* charged Higgs exchange contributions that can compete with the dominant  $W$ -exchange. Experimental measurements of  $B^\pm \rightarrow \tau^\pm\nu_\tau$  [182] initially suggested an enhanced rate with respect to the Standard Model, although the latest results of the Belle Collaboration are consistent with Standard Model expectations. The BaBar Collaboration measured values of the rates for  $\bar{B} \rightarrow D\tau^-\bar{\nu}_\tau$  and  $\bar{B} \rightarrow D^*\tau^-\bar{\nu}_\tau$  [183] that showed a combined  $3.4\sigma$  discrepancy from the Standard Model predictions, which was also not compatible with the Type-II Higgs Yukawa couplings employed by the MSSM. Although subsequent measurements of the Belle and LHCb Collaborations [184] are consistent with the BaBar measurements, the most recent Belle measurements are also compatible (at the  $2\sigma$  level) with either the Standard Model or a Type-II two-Higgs doublet model.

In summary, there are a few hints of possible deviations from the Standard Model in rare  $B$  decays, although none of the discrepancies are significant enough to definitively rule out the Standard Model. The absence of a *significant* deviation in these  $B$ -physics observables from their Standard Model predictions also places useful constraints on the MSSM parameter space [122,157,185].

Finally, we note that the constraints from precision electroweak observables [186] are easily accommodated in models of weak-scale supersymmetry [187]. Thus, robust regions of the MSSM parameter space, compatible with the results of direct and indirect searches for supersymmetry, remain unconstrained.

### I.8. Massive neutrinos in weak-scale supersymmetry:

In the minimal Standard Model and its supersymmetric extension, there are no right-handed neutrinos, and Majorana mass terms for the left-handed neutrinos are absent. However, given the overwhelming evidence for neutrino masses and mixing [188,189], any viable model of fundamental particles must provide a mechanism for generating neutrino masses [190]. In extended supersymmetric models, various mechanisms exist for producing massive neutrinos [191]. Although one can devise models for generating massive Dirac neutrinos [192], the most common approaches for incorporating neutrino masses are based on  $L$ -violating supersymmetric extensions of the MSSM, which generate massive Majorana neutrinos. Two classes of  $L$ -violating supersymmetric models will now be considered.

**I.8.1. The supersymmetric seesaw:** Neutrino masses can be incorporated into the Standard Model by introducing  $SU(3) \times SU(2) \times U(1)$  singlet right-handed neutrinos ( $\nu_R$ ) and super-heavy Majorana masses (typically near the grand unification mass scale) for  $\nu_R$ . In addition, one must also include a standard Yukawa couplings between the lepton doublets, the Higgs doublet, and  $\nu_R$ . The Higgs vacuum expectation value then induces an off-diagonal  $\nu_L$ - $\nu_R$  mass on the order of the electroweak scale. Diagonalizing the neutrino mass matrix (in the three-generation model) yields three superheavy neutrino states, and three very light neutrino states that are identified with the light neutrinos observed in nature. This is the seesaw mechanism [193].

It is straightforward to construct a supersymmetric generalization of the seesaw model of neutrino masses [194,195] by promoting the right-handed neutrino field to a superfield  $\tilde{N}^c = (\tilde{\nu}_R; \nu_R)$ . Integrating out the heavy right-handed neutrino supermultiplet yields a new term in the superpotential [cf. Eq. (1)] of the form

$$W_{\text{seesaw}} = \frac{f}{M_R} (\hat{H}_U \hat{L}) (\hat{H}_U \hat{L}), \quad (22)$$

where  $M_R$  is the mass scale of the right-handed neutrino sector and  $f$  is a dimensionless constant. Note that lepton number is broken by two units, which implies that R-parity is conserved. The supersymmetric analogue of the Majorana neutrino mass term in the sneutrino sector leads to sneutrino-antisneutrino mixing phenomena [195,196].

**I.8.2. R-parity-violating supersymmetry:** In order to incorporate massive neutrinos in renormalizable supersymmetric models while retaining the minimal particle content of the MSSM, one must relax the assumption of R-parity invariance. The most general R-parity-violating (RPV) model involving the MSSM spectrum introduces many new parameters to both the supersymmetry-conserving and the supersymmetry-breaking sectors [197]. Each new interaction term violates either  $B$  or  $L$  conservation. For example, starting from the MSSM superpotential given in Eq. (1) [suitably generalized to

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three generations of quarks, leptons and their superpartners], consider the effect of adding the following new terms:

$$W_{\text{RPV}} = (\lambda_L)_{pmn} \hat{L}_p \hat{L}_m \hat{E}_n^c + (\lambda'_L)_{pmn} \hat{L}_p \hat{Q}_m \hat{D}_n^c + (\lambda_B)_{pmn} \hat{U}_p^c \hat{D}_m^c \hat{D}_n^c + (\mu_L)_p \hat{H}_u \hat{L}_p, \quad (23)$$

where  $p$ ,  $m$ , and  $n$  are generation indices, and gauge group indices are suppressed. Eq. (23) yields new scalar-fermion Yukawa couplings consisting of all possible combinations involving two Standard Model fermions and one scalar superpartner.

Note that the term in Eq. (23) proportional to  $\lambda_B$  violates  $B$ , while the other three terms violate  $L$ . The  $L$ -violating term in Eq. (23) proportional to  $\mu_L$  is the RPV generalization of the  $\mu \hat{H}_u \hat{H}_d$  term of the MSSM superpotential, in which the  $Y = -1$  Higgs/higgsino supermultiplet  $\hat{H}_d$  is replaced by the slepton/lepton supermultiplet  $\hat{L}_p$ .

Phenomenological constraints derived from data on various low-energy  $B$ - and  $L$ -violating processes can be used to establish limits on each of the coefficients  $(\lambda_L)_{pmn}$ ,  $(\lambda'_L)_{pmn}$ , and  $(\lambda_B)_{pmn}$  taken one at a time [197,198]. If more than one coefficient is simultaneously non-zero, then the limits are in general more complicated [199]. All possible RPV terms cannot be simultaneously present and unsuppressed; otherwise the proton decay rate would be many orders of magnitude larger than the present experimental bound. One way to avoid proton decay is to impose  $B$  or  $L$  invariance (either one alone would suffice). Otherwise, one must accept the requirement that certain RPV coefficients must be extremely suppressed.

One particularly interesting class of RPV models is one in which  $B$  is conserved, but  $L$  is violated. It is possible to enforce baryon number conservation (and the stability of the proton), while allowing for lepton-number-violating interactions by imposing a discrete  $\mathbf{Z}_3$  baryon *triality* symmetry on the low-energy theory [200], in place of the standard  $\mathbf{Z}_2$  R-parity. Since the distinction between the Higgs and matter supermultiplets is lost in RPV models where  $L$  is violated, the mixing of sleptons and Higgs bosons, the mixing of neutrinos and neutralinos, and the mixing of charged leptons and charginos are now possible, leading to more complicated mass matrices and mass eigenstates than in the MSSM. Incorporating neutrino masses and mixing in this framework can be found, *e.g.*, in Ref. 201.

Alternatively, one can consider imposing a lepton parity such that all lepton superfields are odd [202,203]. In this case, only the  $B$ -violating term in Eq. (23) survives, and  $L$  is conserved. Models of this type have been considered in Ref. 204. Since  $L$  is conserved in these models, the mixing of the lepton and Higgs superfields is forbidden. However, one expects that lepton parity cannot be exact due to quantum gravity effects. Remarkably, the standard  $\mathbf{Z}_2$  R-parity and the  $\mathbf{Z}_3$  baryon triality are stable with respect to quantum gravity effects, as they can be identified as residual discrete symmetries that arise from broken non-anomalous gauge symmetries [202].

The supersymmetric phenomenology of the RPV models exhibits features that are distinct from that of the MSSM [197].

The LSP is no longer stable, which implies that not all supersymmetric decay chains must yield missing-energy events at colliders. Indeed, the sparticle mass bounds obtained in searches for R-parity-conserving supersymmetry can be considerably relaxed in certain RPV models due to the absence of large missing transverse energy signatures [205]. This can alleviate some of the tension with naturalness discussed in Section I.7.1.

Nevertheless, the loss of the missing-energy signature is often compensated by other striking signals (which depend on which R-parity-violating parameters are dominant). For example, supersymmetric particles in RPV models can be singly produced (in contrast to R-parity-conserving models where supersymmetric particles must be produced in pairs). The phenomenology of pair-produced supersymmetric particles is also modified in RPV models due to new decay chains not present in R-parity-conserving supersymmetry models [197].

In RPV models with lepton number violation (these include weak-scale supersymmetry models with baryon triality mentioned above), both  $\Delta L = 1$  and  $\Delta L = 2$  phenomena are allowed, leading to neutrino masses and mixing [206], neutrinoless double-beta decay [207], sneutrino-antisneutrino mixing [208], and resonant  $s$ -channel production of sneutrinos in  $e^+e^-$  collisions [209] and charged sleptons in  $p\bar{p}$  and  $pp$  collisions [210].

**I.9. Extensions beyond the MSSM:** Extensions of the MSSM have been proposed to solve a variety of theoretical problems. One such problem involves the  $\mu$  parameter of the MSSM. Although  $\mu$  is a supersymmetry-preserving parameter, it must be of order the effective supersymmetry-breaking scale of the MSSM to yield a consistent supersymmetric phenomenology [211]. Any natural solution to the so-called  $\mu$ -problem must incorporate a symmetry that enforces  $\mu = 0$  and a small symmetry-breaking parameter that generates a value of  $\mu$  that is not parametrically larger than the effective supersymmetry-breaking scale [212]. A number of proposed mechanisms in the literature (*e.g.*, see Refs. [211–214]) provide concrete examples of a natural solution to the  $\mu$ -problem of the MSSM.

In extensions of the MSSM, new compelling solutions to the  $\mu$ -problem are possible. For example, one can replace  $\mu$  by the vacuum expectation value of a new  $\text{SU}(3) \times \text{SU}(2) \times \text{U}(1)$  singlet scalar field. This is the NMSSM, which yields phenomena that were briefly discussed in Sections I.4–I.7. The NMSSM superpotential consists only of trilinear terms whose coefficients are dimensionless. There are some advantages to extending the NMSSM further to the USSM [102] by adding a new broken  $\text{U}(1)$  gauge symmetry [215], under which the singlet field is charged.

Alternatively, one can consider a generalized version of the NMSSM (called the GNMSSM in Ref. 170), where all possible renormalizable terms in the superpotential are allowed, which yields new supersymmetric mass terms (analogous to the  $\mu$  term of the MSSM). Although the GNMSSM does not solve the  $\mu$ -problem, it does exhibit regions of parameter space in which the degree of fine-tuning is relaxed, as discussed in Section I.7.1.



The generation of the  $\mu$  term may be connected with the solution to the strong CP problem [216]. Models of this type, which include new gauge singlet fields that are charged under the Peccei-Quinn (PQ) symmetry [217], were first proposed in Ref. 211. The breaking of the PQ symmetry is thus intimately tied to supersymmetry breaking, while naturally yielding a value of  $\mu$  that is of order the electroweak symmetry breaking scale [218].

It is also possible to add higher dimensional Higgs multiplets, such as Higgs triplet superfields [219], provided a custodial-symmetric model (in which the  $\rho$ -parameter of precision electroweak physics is close to 1 [186]) can be formulated. Such models can provide a rich phenomenology of new signals for future LHC studies.

All supersymmetric models discussed in this review possess self-conjugate fermions—the Majorana gluinos and neutralinos. However, it is possible to add additional chiral superfields in the adjoint representation. The spin-1/2 components of these new superfields can pair up with the gauginos to form Dirac gauginos [220,221]. Such states appear in models of so-called supersoft supersymmetry breaking [222], in some generalized GMSB models [223] and in R-symmetric supersymmetry [29,224]. Such approaches often lead to improved naturalness and/or significantly relaxed flavor constraints. The implications of models of Dirac gauginos on the observed Higgs boson mass and its properties is addressed in Ref. 225.

For completeness, we briefly note other MSSM extensions considered in the literature. These include an enlarged electroweak gauge group beyond  $SU(2) \times U(1)$  [226]; and/or the addition of new (possibly exotic) matter supermultiplets such as vector-like fermions and their superpartners [171,227].

## References

1. The early history of supersymmetry and a guide to the original literature can be found in *The Supersymmetric World—The Beginnings of the Theory*, edited by G. Kane and M. Shifman (World Scientific, Singapore, 2000).
2. R. Haag, J.T. Lopuszanski, and M. Sohnius, Nucl. Phys. **B88**, 257 (1975); S.R. Coleman and J. Mandula, Phys. Rev. **159**, 1251 (1967).
3. H.P. Nilles, Phys. Reports **110**, 1 (1984).
4. P. Nath, R. Arnowitt, and A.H. Chamseddine, *Applied  $N = 1$  Supergravity* (World Scientific, Singapore, 1984).
5. S. Weinberg, *The Quantum Theory of Fields, Volume III: Supersymmetry* (Cambridge University Press, Cambridge, UK, 2000); P. Binétruy, *Supersymmetry: Theory, Experiment, and Cosmology* (Oxford University Press, Oxford, UK, 2006).
6. S.P. Martin, in *Perspectives on Supersymmetry II*, edited by G.L. Kane (World Scientific, Singapore, 2010) pp. 1–153; see <http://www.niu.edu/spmartin/primer/> for the latest version and errata.
7. E. Witten, Nucl. Phys. **B188**, 513 (1981).
8. S. Dimopoulos and H. Georgi, Nucl. Phys. **B193**, 150 (1981).
9. N. Sakai, Z. Phys. **C11**, 153 (1981); R.K. Kaul, Phys. Lett. **109B**, 19 (1982); R.K. Kaul and M. Parthasarathi, Nucl. Phys. **B199**, 36 (1982).
10. L. Susskind, Phys. Reports **104**, 181 (1984).
11. L. Girardello and M. Grisaru, Nucl. Phys. **B194**, 65 (1982).
12. L.J. Hall and L. Randall, Phys. Rev. Lett. **65**, 2939 (1990); I. Jack and D.R.T. Jones, Phys. Lett. **B457**, 101 (1999).
13. O. Buchmüller and P. de Jong, “*Supersymmetry Part II (Experiment)*,” and associated *Particle Listings: Other Searches—Supersymmetric Particles*, in the web edition of the *Review of Particle Physics* at <http://pdg.lbl.gov>.
14. V.F. Weisskopf, Phys. Rev. **56**, 72 (1939).
15. For a review, see N. Polonsky, *Supersymmetry: Structure and phenomena. Extensions of the standard model*, Lect. Notes Phys. **M68**, 1 (2001).
16. G. Bertone, D. Hooper, and J. Silk, Phys. Reports **405**, 279 (2005).
17. D. Hooper, “TASI 2008 Lectures on Dark Matter,” in *The Dawn of the LHC Era, Proceedings of the 2008 Theoretical and Advanced Study Institute in Elementary Particle Physics*, Boulder, Colorado, 2–27 June 2008, edited by Tao Han (World Scientific, Singapore, 2009).
18. H. Pagels and J.R. Primack, Phys. Rev. Lett. **48**, 223 (1982).
19. H. Goldberg, Phys. Rev. Lett. **50**, 1419 (1983) [Erratum: **103**, 099905 (2009)].
20. J. Ellis *et al.*, Nucl. Phys. **B238**, 453 (1984).
21. G. Jungman, M. Kamionkowski, and K. Griest, Phys. Reports **267**, 195 (1996); K. Griest and M. Kamionkowski, Phys. Reports **333**, 167 (2000).
22. F.D. Steffen, Eur. Phys. J. **C59**, 557 (2009).
23. M. Drees and G. Gerbier, “*Dark Matter*,” in the web edition of the *Review of Particle Physics* at <http://pdg.lbl.gov>.
24. M. Dine and A. Kusenko, Rev. Mod. Phys. **76**, 1 (2003).
25. For a review see e.g., J.M. Cline, in *Particle Physics and Cosmology: the Fabric of Spacetime*, edited by F. Bernardeau, C. Grojean, and J. Dalibard, Proceedings of Les Houches Summer School **86**, 53–116 (2007).
26. A.D. Sakharov, Sov. Phys. JETP Lett. **5**, 24 (1967); Sov. Phys. JETP **49**, 594 (1979).
27. Reviews of electroweak baryogenesis can be found in A.G. Cohen, D. Kaplan, and A. Nelson, Ann. Rev. Nucl. and Part. Sci. **43**, 27 (1993); D.E. Morrissey and M.J. Ramsey-Musolf, New J. Phys. **14**, 125003 (2012).
28. T. Cohen, D.E. Morrissey, and A. Pierce, Phys. Rev. **D86**, 013009 (2012); D. Curtin, P. Jaiswal, and P. Meade, JHEP **1208**, 005 (2012); M. Carena *et al.*, JHEP **1302**, 001 (2013).
29. U. Sarkar and R. Adhikari, Phys. Rev. **D55**, 3836 (1997); R. Fok *et al.*, Phys. Rev. **D87**, 055018 (2013).
30. I. Affleck and M. Dine, Nucl. Phys. **B249**, 361 (1985).
31. See e.g., K. Hamaguchi, hep-ph/0212305 (2002).
32. H.E. Haber and G.L. Kane, Phys. Reports **117**, 75 (1985).

# Searches Particle Listings

## Supersymmetric Particle Searches

33. M. Drees, R. Godbole, and P. Roy, *Theory and Phenomenology of Sparticles* (World Scientific, Singapore, 2005);  
H. Baer and X. Tata, *Weak Scale Supersymmetry: from Superfields to Scattering Events* (Cambridge University Press, Cambridge, UK, 2006);  
I.J.R. Aitchison, *Supersymmetry in Particle Physics: an elementary introduction* (Cambridge University Press, Cambridge, UK, 2007).
34. Our notation for the charge-conjugated fields follows the notation of P. Langacker, *The Standard Model and Beyond* (CRC Press, Boca Raton, FL, 2010).
35. P. Fayet, Nucl. Phys. **B90**, 104 (1975).
36. K. Inoue *et al.*, Prog. Theor. Phys. **67**, 1889 (1982);  
R. Flores and M. Sher, Ann. Phys. (NY) **148**, 95 (1983).
37. J.F. Gunion and H.E. Haber, Nucl. Phys. **B272**, 1 (1986) [Erratum: **B402**, 567 (1993)].
38. J. Wess and J. Bagger, *Supersymmetry and Supergravity* (Princeton University Press, Princeton, NJ, 1992).
39. I. Buchbinder, S. Kuzenko, and J. Yarevskaya, Nucl. Phys. **B411**, 665 (1994);  
I. Antoniadis, E. Dudas, and D.M. Ghilencea, JHEP **0803**, 045 (2008).
40. For an overview of the theory and models of the soft-supersymmetry-breaking Lagrangian, see D.J.H. Chung *et al.*, Phys. Reports **407**, 1 (2005).
41. J. Ellis *et al.*, Phys. Rev. **D39**, 844 (1989);  
U. Ellwanger and C. Hugonie, Eur. Phys. J. **C25**, 297 (2002);  
U. Ellwanger, C. Hugonie, and A.M. Teixeira, Phys. Reports **496**, 1 (2010);  
M. Maniatis, Int. J. Mod. Phys. **A25**, 3505 (2010).
42. P. Fayet, Phys. Lett. **69B**, 489 (1977);  
G. Farrar and P. Fayet, Phys. Lett. **76B**, 575 (1978).
43. P. Fayet, Phys. Lett. **84B**, 421 (1979); Phys. Lett. **86B**, 272 (1979).
44. D.Z. Freedman and A. Van Proeyen, *Supergravity* (Cambridge University Press, Cambridge, UK, 2012).
45. S. Deser and B. Zumino, Phys. Rev. Lett. **38**, 1433 (1977);  
E. Cremmer *et al.*, Phys. Lett. **79B**, 231 (1978).
46. R. Casalbuoni *et al.*, Phys. Lett. **B215**, 313 (1988); Phys. Rev. **D39**, 2281 (1989);  
A.L. Maroto and J.R. Pelaez, Phys. Rev. **D62**, 023518 (2000).
47. Z. Komargodski and N. Seiberg, JHEP **0909**, 066 (2009);  
I. Antoniadis *et al.*, Theor. Math. Phys. **170**, 26 (2012).
48. K. Benakli and C. Moura, Nucl. Phys. **B791**, 125 (2008);  
C. Cheung, Y. Nomura, and J. Thaler, JHEP **1003**, 073 (2010);  
N. Craig, J. March-Russell, and M. McCullough, JHEP **1010**, 095 (2010);  
R. Argurio, Z. Komargodski and A. Mariotti, Phys. Rev. Lett. **107**, 061601 (2011).
49. A.H. Chamseddine, R. Arnowitt, and P. Nath, Phys. Rev. Lett. **49**, 970 (1982);  
R. Barbieri, S. Ferrara, and C.A. Savoy, Phys. Lett. **119B**, 343 (1982);  
L. Ibáñez, Nucl. Phys. **B218**, 514 (1982);  
H.-P. Nilles, M. Srednicki, and D. Wyler, Phys. Lett. **120B**, 346 (1983); Phys. Lett. **124B**, 337 (1983);  
E. Cremmer, P. Fayet, and L. Girardello, Phys. Lett. **122B**, 41 (1983);  
N. Ohta, Prog. Theor. Phys. **70**, 542 (1983).
50. L. Alvarez-Gaumé, J. Polchinski, and M.B. Wise, Nucl. Phys. **B221**, 495 (1983).
51. L.J. Hall, J. Lykken, and S. Weinberg, Phys. Rev. **D27**, 2359 (1983).
52. S.K. Soni and H.A. Weldon, Phys. Lett. **126B**, 215 (1983);  
Y. Kawamura, H. Murayama, and M. Yamaguchi, Phys. Rev. **D51**, 1337 (1995).
53. See, *e.g.*, A. Brignole, L.E. Ibáñez, and C. Muñoz, in *Perspectives on Supersymmetry II*, edited by G.L. Kane (World Scientific, Singapore, 2010) pp. 244–268.
54. A.B. Lahanas and D.V. Nanopoulos, Phys. Reports **145**, 1 (1987).
55. J.L. Feng, A. Rajaraman, and F. Takayama, Phys. Rev. Lett. **91**, 011302 (2003); Phys. Rev. **D68**, 063504 (2003);  
Gen. Rel. Grav. **36**, 2575 (2004).
56. L. Randall and R. Sundrum, Nucl. Phys. **B557**, 79 (1999).
57. F. D’Eramo, J. Thaler, and Z. Thomas, JHEP **1206**, 151 (2012); JHEP **1309**, 125 (2013);  
S.P. de Alwis, Phys. Rev. **D77**, 105020 (2008); JHEP **1301**, 006 (2013);  
K. Harigaya and M. Ibe, Phys. Rev. **D90**, 085028 (2014).
58. See *e.g.*, I. Jack, D.R.T. Jones, and R. Wild, Phys. Lett. **B535**, 193 (2002);  
B. Murakami and J.D. Wells, Phys. Rev. **D68**, 035006 (2003);  
R. Kitano, G.D. Kribs, and H. Murayama, Phys. Rev. **D70**, 035001 (2004);  
R. Hodgson *et al.*, Nucl. Phys. **B728**, 192 (2005);  
D.R.T. Jones and G.G. Ross, Phys. Lett. **B642**, 540 (2006).
59. S. Asai *et al.*, Phys. Lett. **B653**, 81 (2007).
60. M. Dine, W. Fischler, and M. Srednicki, Nucl. Phys. **B189**, 575 (1981);  
S. Dimopoulos and S. Raby, Nucl. Phys. **B192**, 353 (1982); Nucl. Phys. **B219**, 479 (1983);  
M. Dine and W. Fischler, Phys. Lett. **110B**, 227 (1982);  
C. Nappi and B. Ovrut, Phys. Lett. **113B**, 175 (1982);  
L. Alvarez-Gaumé, M. Claudson, and M. Wise, Nucl. Phys. **B207**, 96 (1982).
61. M. Dine and A.E. Nelson, Phys. Rev. **D48**, 1277 (1993);  
M. Dine, A.E. Nelson, and Y. Shirman, Phys. Rev. **D51**, 1362 (1995);  
M. Dine *et al.*, Phys. Rev. **D53**, 2658 (1996).
62. G.F. Giudice and R. Rattazzi, Phys. Reports **322**, 419 (1999).
63. E. Poppitz and S.P. Trivedi, Phys. Rev. **D55**, 5508 (1997);  
H. Murayama, Phys. Rev. Lett. **79**, 18 (1997);  
M.A. Luty and J. Terning, Phys. Rev. **D57**, 6799 (1998);  
K. Agashe, Phys. Lett. **B435**, 83 (1998);  
N. Arkani-Hamed, J. March-Russell, and H. Murayama, Nucl. Phys. **B509**, 3 (1998);  
C. Csaki, Y. Shirman, and J. Terning, JHEP **0705**, 099 (2007);  
M. Ibe and R. Kitano, Phys. Rev. **D77**, 075003 (2008).
64. M. Kawasaki *et al.*, Phys. Rev. **D78**, 065011 (2008).

65. M.J. Strassler and K.M. Zurek, Phys. Lett. **B651**, 374 (2007);  
T. Han *et al.*, JHEP **0807**, 008 (2008).
66. M.J. Strassler, hep-ph/0607160;  
K.M. Zurek, Phys. Rev. **D79**, 115002 (2009).
67. See *e.g.*, M. Quiros, in *Particle Physics and Cosmology: The Quest for Physics Beyond the Standard Model(s), Proceedings of the 2002 Theoretical Advanced Study Institute in Elementary Particle Physics (TASI 2002)*, edited by H.E. Haber and A.E. Nelson (World Scientific, Singapore, 2004) pp. 549–601;  
C. Csaki, in *ibid.*, pp. 605–698.
68. See, *e.g.*, J. Parsons and A. Pomarol, “Extra Dimensions,” in the web edition of the *Review of Particle Physics* at <http://pdg.lbl.gov>.
69. See *e.g.*, V.A. Rubakov, Sov. Phys. Usp. **44**, 871 (2001);  
J. Hewett and M. Spiropulu, Ann. Rev. Nucl. and Part. Sci. **52**, 397 (2002).
70. Z. Chacko, M.A. Luty, and E. Ponton, JHEP **0007**, 036 (2000);  
D.E. Kaplan, G.D. Kribs, and M. Schmaltz, Phys. Rev. **D62**, 035010 (2000);  
Z. Chacko *et al.*, JHEP **0001**, 003 (2000).
71. J. Scherk and J.H. Schwarz, Phys. Lett. **82B**, 60 (1979);  
Nucl. Phys. **B153**, 61 (1979).
72. See, *e.g.*, R. Barbieri, L.J. Hall, and Y. Nomura, Phys. Rev. **D66**, 045025 (2002); Nucl. Phys. **B624**, 63 (2002).
73. J. Wells, hep-ph/0306127; Phys. Rev. **D71**, 015013 (2005).
74. N. Arkani-Hamed and S. Dimopoulos, JHEP **0506**, 073 (2005);  
G.F. Giudice and A. Romanino, Nucl. Phys. **B699**, 65 (2004) [Erratum: **B706**, 487 (2005)].
75. G. Aad *et al.* [ATLAS and CMS Collab.] Phys. Rev. Lett. **114**, 191803 (2015).
76. G.F. Giudice and A. Strumia, Nucl. Phys. **B858**, 63 (2012).
77. A. Arvanitaki *et al.*, JHEP **1302**, 126 (2013);  
N. Arkani-Hamed *et al.*, arXiv:1212.6971(2012).
78. J.P. Vega and G. Villadoro, JHEP **1507**, 159 (2015).
79. Y. Kahn, M. McCullough and J. Thaler, JHEP **1311**, 161 (2013).
80. L.J. Hall and Y. Nomura, JHEP **1201**, 082 (2012);  
M. Ibe and T.T. Yanagida, Phys. Lett. **B709**, 374 (2012).
81. H.E. Haber, in *Recent Directions in Particle Theory, Proceedings of the 1992 Theoretical Advanced Study Institute in Particle Physics*, edited by J. Harvey and J. Polchinski (World Scientific, Singapore, 1993) pp. 589–686.
82. J.M. Frere, D.R.T. Jones, and S. Raby, Nucl. Phys. **B222**, 11 (1983);  
J.P. Derendinger and C.A. Savoy, Nucl. Phys. **B237**, 307 (1984);  
J.F. Gunion, H.E. Haber, and M. Sher, Nucl. Phys. **B306**, 1 (1988);  
J.A. Casas, A. Lleyda, and C. Munoz, Nucl. Phys. **B471**, 3 (1996).
83. S.P. Martin, Phys. Rev. **D61**, 035004 (2000).
84. S. Dimopoulos and D. Sutter, Nucl. Phys. **B452**, 496 (1995);  
D.W. Sutter, Stanford Ph. D. thesis, hep-ph/9704390.
85. H.E. Haber, Nucl. Phys. B (Proc. Suppl.) **62A-C**, 469 (1998).
86. R.M. Barnett, J.F. Gunion, and H.E. Haber, Phys. Lett. **B315**, 349 (1993);  
H. Baer, X. Tata, and J. Woodside, Phys. Rev. **D41**, 906 (1990).
87. S.M. Bilenky, E.Kh. Khristova, and N.P. Nedelcheva, Phys. Lett. **B161**, 397 (1985); Bulg. J. Phys. **13**, 283 (1986);  
G. Moortgat-Pick and H. Fraas, Eur. Phys. J. **C25**, 189 (2002).
88. J. Rosiek, Phys. Rev. **D41**, 3464 (1990) [Erratum: hep-ph/9511250]. The most recent corrected version of this manuscript can be found on the author’s webpage, <http://www.fuw.edu.pl/~rosiek/physics/prd41.html>.
89. J. Alwall *et al.*, JHEP **0709**, 028 (2007). See also the MadGraph homepage, <http://madgraph.hep.uiuc.edu>.
90. T. Hahn, Comp. Phys. Comm. **140**, 418 (2001);  
T. Hahn and C. Schappacher, Comp. Phys. Comm. **143**, 54 (2002). The FeynArts homepage is located at <http://www.feynarts.de>.
91. A. Pukhov *et al.*, INP MSU report 98-41/542 (hep-ph/9908288);  
E. Boos *et al.* [CompHEP Collab.], Nucl. Instrum. Methods **A534**, 50 (2004). The CompHEP homepage is located at <http://comphep.sinp.msu.ru>.
92. D.M. Pierce *et al.*, Nucl. Phys. **B491**, 3 (1997).
93. P. Skands *et al.*, JHEP **0407**, 036 (2004);  
B.C. Allanach *et al.*, Comp. Phys. Comm. **180**, 8 (2009). The Supersymmetry Les Houches Accord homepage is <http://skands.physics.monash.edu/slha/>.
94. For further details, see *e.g.*, Appendix C of Ref. 32 and Appendix A of Ref. 37.
95. J.L. Kneur and G. Moultaka, Phys. Rev. **D59**, 015005 (1999).
96. S.Y. Choi *et al.*, Eur. Phys. J. **C14**, 535 (2000).
97. R.A. Horn and C.R. Johnson, *Matrix Analysis*, 2nd Edition (Cambridge University Press, Cambridge, UK, 2003).
98. H.K. Dreiner, H.E. Haber, and S.P. Martin, Phys. Reports **494**, 1 (2010).
99. S.Y. Choi *et al.*, Eur. Phys. J. **C22**, 563 (2001); Eur. Phys. J. **C23**, 769 (2002).
100. G.J. Gounaris, C. Le Mouel, and P.I. Porfyriadis, Phys. Rev. **D65**, 035002 (2002);  
G.J. Gounaris and C. Le Mouel, Phys. Rev. **D66**, 055007 (2002).
101. T. Takagi, Japan J. Math. **1**, 83 (1925).
102. S.Y. Choi *et al.*, Nucl. Phys. **B778**, 85 (2007).
103. M.M. El Kheishen, A.A. Aboshousha, and A.A. Shafik, Phys. Rev. **D45**, 4345 (1992);  
M. Guchait, Z. Phys. **C57**, 157 (1993) [Erratum: **C61**, 178 (1994)].
104. T. Hahn, preprint MPP-2006-85, physics/0607103.
105. K. Hikasa and M. Kobayashi, Phys. Rev. **D36**, 724 (1987);  
F. Gabbiani and A. Masiero, Nucl. Phys. **B322**, 235 (1989);  
Ph. Brax and C.A. Savoy, Nucl. Phys. **B447**, 227 (1995).
106. J. Ellis and S. Rudaz, Phys. Lett. **128B**, 248 (1983);  
F. Browning, D. Chang, and W.Y. Keung, Phys. Rev.

# Searches Particle Listings

## Supersymmetric Particle Searches

- D64**, 015010 (2001);  
A. Bartl *et al.*, Phys. Lett. **B573**, 153 (2003); Phys. Rev. **D70**, 035003 (2004).
107. J.F. Gunion *et al.*, *The Higgs Hunter's Guide* (Westview Press, Boulder, CO, 2000);  
M. Carena and H.E. Haber, Prog. in Part. Nucl. Phys. **50**, 63 (2003);  
A. Djouadi, Phys. Reports **459**, 1 (2008).
  108. P. Bechtle *et al.*, Eur. Phys. J. **C73**, 2354 (2013);  
M. Carena *et al.*, Eur. Phys. J. **C73**, 2552 (2013).
  109. H.E. Haber and M. Sher, Phys. Rev. **D35**, 2206 (1987).
  110. L.J. Hall, D. Pinner, and J.T. Ruderman, JHEP **1204**, 131 (2012).
  111. L.J. Hall and M.B. Wise, Nucl. Phys. **B187**, 397 (1981).
  112. H.E. Haber and R. Hempfling, Phys. Rev. Lett. **66**, 1815 (1991);  
Y. Okada, M. Yamaguchi, and T. Yanagida, Prog. Theor. Phys. **85**, 1 (1991);  
J. Ellis, G. Ridolfi, and F. Zwirner, Phys. Lett. **B257**, 83 (1991).
  113. See, *e.g.*, G. Degrandi *et al.*, Eur. Phys. J. **C28**, 133 (2003);  
S.P. Martin, Phys. Rev. **D75**, 055005 (2007);  
P. Kant *et al.*, JHEP **1008**, 104 (2010);  
T. Hahn *et al.*, Phys. Rev. Lett. **112**, 141801 (2014);  
J.P. Vega and G. Villadoro, JHEP **1507**, 159 (2015).
  114. A. Pilaftsis and C.E.M. Wagner, Nucl. Phys. **B553**, 3 (1999);  
D.A. Demir, Phys. Rev. **D60**, 055006 (1999);  
S.Y. Choi, M. Drees, and J.S. Lee, Phys. Lett. **B481**, 57 (2000);  
M. Carena *et al.*, Nucl. Phys. **B586**, 92 (2000); Phys. Lett. **B495**, 155 (2000); Nucl. Phys. **B625**, 345 (2002);  
M. Frank *et al.*, JHEP **0702**, 047 (2007);  
S. Heinemeyer *et al.*, Phys. Lett. **B652**, 300 (2007).
  115. H.E. Haber and J.D. Mason, Phys. Rev. **D77**, 115011 (2008).
  116. S. Khalil, Int. J. Mod. Phys. **A18**, 1697 (2003).
  117. W. Fischler, S. Paban, and S. Thomas, Phys. Lett. **B289**, 373 (1992);  
S.M. Barr, Int. J. Mod. Phys. **A8**, 209 (1993);  
T. Ibrahim and P. Nath, Phys. Rev. **D58**, 111301 (1998) [Erratum: **D60**, 099902 (1999)];  
M. Brhlik, G.J. Good, and G.L. Kane, Phys. Rev. **D59**, 115004 (1999);  
V.D. Barger *et al.*, Phys. Rev. **D64**, 056007 (2001);  
S. Abel, S. Khalil, and O. Lebedev, Nucl. Phys. **B606**, 151 (2001);  
K.A. Olive *et al.*, Phys. Rev. **D72**, 075001 (2005);  
G.F. Giudice and A. Romanino, Phys. Lett. **B634**, 307 (2006).
  118. A. Masiero and L. Silvestrini, in *Perspectives on Supersymmetry*, edited by G.L. Kane (World Scientific, Singapore, 1998) pp. 423–441.
  119. M. Pospelov and A. Ritz, Ann. Phys. **318**, 119 (2005).
  120. See, *e.g.*, F. Gabbiani *et al.*, Nucl. Phys. **B477**, 321 (1996);  
A. Masiero, and O. Vives, New J. Phys. **4**, 1 (2002).
  121. For a review and references to the original literature, see: M.J. Ramsey-Musolf and S. Su, Phys. Reports **456**, 1 (2008).
  122. M. Carena, A. Menon, and C.E.M. Wagner, Phys. Rev. **D79**, 075025 (2009);  
S. Jager, Eur. Phys. J. **C59**, 497 (2009);  
W. Altmannshofer *et al.*, Nucl. Phys. **B830**, 17 (2010).
  123. M.B. Einhorn and D.R.T. Jones, Nucl. Phys. **B196**, 475 (1982).
  124. For a review, see R.N. Mohapatra, in *Particle Physics 1999*, ICTP Summer School in Particle Physics, Trieste, Italy, edited by G. Senjanovic and A.Yu. Smirnov (World Scientific, Singapore, 2000) pp. 336–394;  
W.J. Marciano and G. Senjanovic, Phys. Rev. **D25**, 3092 (1982).
  125. L.E. Ibáñez and G.G. Ross, Phys. Lett. **B110**, 215 (1982).
  126. J. Abdullah *et al.* [DELPHI Collab.], Eur. Phys. J. **C31**, 421 (2004).
  127. H.K. Dreiner *et al.*, Eur. Phys. J. **C62**, 547 (2009).
  128. G.F. Giudice *et al.*, JHEP **9812**, 027 (1998);  
A. Pomarol and R. Rattazzi, JHEP **9905**, 013 (1999);  
D.W. Jung and J.Y. Lee, JHEP **0903**, 123 (2009).
  129. J.F. Gunion and H.E. Haber, Phys. Rev. **D37**, 2515 (1988);  
S.Y. Choi, M. Drees, and B. Gaissmaier, Phys. Rev. **D70**, 014010 (2004).
  130. J.L. Feng *et al.*, Phys. Rev. Lett. **83**, 1731 (1999);  
J.F. Gunion and S. Mrenna, Phys. Rev. **D62**, 015002 (2000).
  131. T. Gherghetta, G.F. Giudice, and J.D. Wells, Nucl. Phys. **B559**, 27 (1999).
  132. M. Endo, M. Yamaguchi, and K. Yoshioka, Phys. Rev. **D72**, 015004 (2005);  
K. Choi, K.S. Jeong, and K.-I. Okumura, JHEP **0509**, 039 (2005);  
O. Loaiza-Brito *et al.*, AIP Conf. Proc. **805**, 198 (2006).
  133. See *e.g.*, G. D'Ambrosio *et al.*, Nucl. Phys. **B465**, 155 (2002).
  134. For a review of minimal flavor violation in supersymmetric theories, see C. Smith, Acta Phys. Polon. Supp. **3**, 53 (2010).
  135. M. Drees and S.P. Martin, in *Electroweak Symmetry Breaking and New Physics at the TeV Scale*, edited by T. Barklow *et al.*, (World Scientific, Singapore, 1996) pp. 146–215.
  136. G.L. Kane *et al.*, Phys. Rev. D **49**, 6173 (1994).
  137. J.R. Ellis *et al.*, Phys. Lett. **B573**, 162 (2003); Phys. Rev. **D70**, 055005 (2004).
  138. L.E. Ibáñez and D. Lust, Nucl. Phys. **B382**, 305 (1992);  
B. de Carlos, J.A. Casas, and C. Muñoz, Phys. Lett. **B299**, 234 (1993);  
V. Kaplunovsky and J. Louis, Phys. Lett. **B306**, 269 (1993);  
A. Brignole, L.E. Ibáñez, and C. Muñoz, Nucl. Phys. **B422**, 125 (1994) [Erratum: **B436**, 747 (1995)].
  139. A. Arbey *et al.*, Phys. Rev. **D87**, 115020 (2013).
  140. P. Draper *et al.*, Phys. Rev. **D85**, 095007 (2012).
  141. P. Meade, N. Seiberg, and D. Shih, Prog. Theor. Phys. Supp. **177**, 143 (2009);  
M. Buican *et al.*, JHEP **0903**, 016 (2009).
  142. A. Rajaraman *et al.*, Phys. Lett. **B678**, 367 (2009);  
L.M. Carpenter *et al.*, Phys. Rev. **D79**, 035002 (2009).

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# Searches Particle Listings

## Supersymmetric Particle Searches

143. S. Ambrosanio, G.D. Kribs, and S.P. Martin, Nucl. Phys. **B516**, 55 (1998).
144. For a review and guide to the literature, see J.F. Gunion and H.E. Haber, in *Perspectives on Supersymmetry II*, edited by G.L. Kane (World Scientific, Singapore, 2010) pp. 420–445.
145. A. de Gouvea, A. Friedland, and H. Murayama, Phys. Rev. **D57**, 5676 (1998).
146. T. Han, D. Marfatia, and R.-J. Zhang, Phys. Rev. **D61**, 013007 (2000);  
Z. Chacko and E. Ponton, Phys. Rev. **D66**, 095004 (2002);  
A. Delgado, G.F. Giudice, and P. Slavich, Phys. Lett. **B653**, 424 (2007);  
T. Liu and C.E.M. Wagner, JHEP **0806**, 073 (2008).
147. B. Allanach *et al.*, Phys. Rev. **D92**, 015006 (2015).
148. A. Djouadi, J.L. Kneur, and G. Moultaka, Comp. Phys. Comm. **176**, 426 (2007);  
C.F. Berger *et al.*, JHEP **0902**, 023 (2009).
149. K.J. de Vries *et al.*, Eur. Phys. J. **C75**, 422 (2015).
150. M. Cahill-Rowley *et al.*, Phys. Rev. **D88**, 035002 (2013);  
Phys. Rev. **D91**, 055002 (2015).
151. N. Arkani-Hamed *et al.*, hep-ph/0703088;  
J. Alwall *et al.*, Phys. Rev. **D79**, 015005 (2009);  
J. Alwall, P. Schuster, and N. Toro, Phys. Rev. **D79**, 075020 (2009);  
D.S.M. Alves, E. Izaguirre, and J.G. Wacker, Phys. Lett. **B702**, 64 (2011); JHEP **1110**, 012 (2011);  
D. Alves *et al.*, J. Phys. **G39**, 105005 (2012).
152. H. Baer *et al.*, JHEP **1205**, 109 (2012).
153. R. Barbieri and G.F. Giudice, Nucl. Phys. **B305**, 63 (1988).
154. G.W. Anderson and D.J. Castano, Phys. Lett. **B347**, 300 (1995); Phys. Rev. **D52**, 1693 (1995); Phys. Rev. **D53**, 2403 (1996);  
J.L. Feng, K.T. Matchev, and T. Moroi, Phys. Rev. **D61**, 075005 (2000);  
P. Athron and D.J. Miller, Phys. Rev. **D76**, 075010 (2007);  
M.E. Cabrera, J.A. Casas, and R.R. de Austri, JHEP **0903**, 075 (2009) (2009);  
H. Baer *et al.*, Phys. Rev. Lett. **109**, 161802 (2012).
155. D.M. Ghilencea and G.G. Ross, Nucl. Phys. **B868**, 65 (2013).
156. G.L. Kane and S.F. King, Phys. Lett. **B451**, 113 (1999);  
M. Bastero-Gil, G.L. Kane, and S.F. King, Phys. Lett. **B474**, 103 (2000);  
J.A. Casas, J.R. Espinosa, and I. Hidalgo, JHEP **0401**, 008 (2004);  
J. Abe, T. Kobayashi, and Y. Omura, Phys. Rev. **D76**, 015002 (2007);  
R. Essig and J.-F. Fortin, JHEP **0804**, 073 (2008).
157. O. Buchmüller *et al.*, Eur. Phys. J. **C74**, 2922 (2014).
158. P. Bechtle *et al.*, Eur. Phys. J. **C76**, 96 (2016).
159. R. Barbieri and A. Strumia, hep-ph/0007265.
160. L. Giusti, A. Romanino, and A. Strumia, Nucl. Phys. **B550**, 3 (1999);  
H.C. Cheng and I. Low, JHEP **0309**, 051 (2003); JHEP **0408**, 061 (2004);  
R. Harnik *et al.*, Phys. Rev. **D70**, 015002 (2004).
161. H. Baer, V. Barger, and D. Mickelson, Phys. Rev. **D88**, 095013 (2013).
162. H. Baer *et al.*, Phys. Rev. **D87**, 035017 (2013); Phys. Rev. **D87**, 115028 (2013);  
Ann. Rev. Nucl. and Part. Sci. **63**, 351 (2013).
163. J. Feng, K. Matchev, and T. Moroi, Phys. Rev. Lett. **84**, 2322 (2000); Phys. Rev. **D61**, 075005 (2000);  
J. Feng and F. Wilczek, Phys. Lett. **B631**, 170 (2005);  
D. Horton and G.G. Ross, Nucl. Phys. **B830**, 221 (2010).
164. M. Drees, Phys. Rev. **D33**, 1468 (1986);  
S. Dimopoulos and G.F. Giudice, Phys. Lett. **B357**, 573 (1995);  
A. Pomarol and D. Tommasini, Nucl. Phys. **B466**, 3 (1996).
165. M. Dine, A. Kagan, and S. Samuel, Phys. Lett. **B243**, 250 (1990);  
A.G. Cohen, D.B. Kaplan, and A.E. Nelson, Phys. Lett. **B388**, 588 (1996).
166. K.L. Chan, U. Chattopadhyay, and P. Nath, Phys. Rev. **D58**, 096004 (1998);  
R. Kitano and Y. Nomura, Phys. Rev. **D73**, 095004 (2006);  
M. Perelstein and C. Spethmann, JHEP **0704**, 070 (2007);  
H. Abe, T. Kobayashi, and Y. Omura, Phys. Rev. **D76**, 015002 (2007);  
D. Horton and G.G. Ross, Nucl. Phys. **B830**, 221 (2010);  
H. Baer *et al.*, JHEP **1010**, 018 (2010);  
M. Asano *et al.*, JHEP **1012**, 019 (2010);  
H. Baer, V. Barger, and P. Huang, JHEP **1111**, 031 (2011);  
C. Brust *et al.*, JHEP **1203**, 103 (2012);  
M. Papucci, J.T. Ruderman, and A. Weiler, JHEP **1209**, 035 (2012);  
H.K. Dreiner, M. Kramer, and J. Tattersall, Europhys. Lett. **99**, 61001 (2012).
167. S.P. Martin, Phys. Rev. **D75**, 115005 (2007); Phys. Rev. **D78**, 055019 (2009).
168. J. Fan, M. Reece, and J.T. Ruderman, JHEP **1111**, 012 (2011); JHEP **1207**, 196 (2012).
169. R. Dermisek and J.F. Gunion, Phys. Rev. Lett. **95**, 041801 (2005); Phys. Rev. **D75**, 095019 (2007); Phys. Rev. **D76**, 095006 (2007).
170. G.G. Ross and K. Schmidt-Hoberg, Nucl. Phys. **B862**, 710 (2012); JHEP **1208**, 074 (2012);  
A. Kaminska, G.G. Ross, and K. Schmidt-Hoberg, JHEP **1311**, 209 (2013).
171. S.P. Martin and J.D. Wells, Phys. Rev. **D86**, 035017 (2012).
172. B. Bellazzini *et al.*, Phys. Rev. **D79**, 095003 (2009).
173. H. Baer, V. Barger, and M. Savoy, Phys. Rev. **D93**, 035016 (2016).
174. D. Stockinger, J. Phys. **G34**, R45 (2007);  
P. Athron *et al.*, Eur. Phys. J. **C76**, 62 (2016).
175. T. Blum *et al.*, arXiv:1311.2198;  
A. E. Dorokhova, A. E. Radzhabov, A.S. Zhevlakov, Sov. Phys. JETP Lett. **100**, 133 (2014).
176. M. Ibe, T. T. Yanagida, and N. Yokozaki, JHEP **1308**, 067 (2013).
177. A. Limosani *et al.* [Belle Collab.], Phys. Rev. Lett. **103**, 241801 (2009);  
J.P. Lees *et al.* [BaBar Collab.], Phys. Rev. Lett. **109**, 191801 (2012); Phys. Rev. **D86**, 112008 (2012).
178. M. Misiak *et al.*, Phys. Rev. Lett. **114**, 221801 (2015).

# Searches Particle Listings

## Supersymmetric Particle Searches

179. See, *e.g.*, M. Ciuchini *et al.*, Phys. Rev. **D67**, 075016 (2003);  
T. Hurth, Rev. Mod. Phys. **75**, 1159 (2003);  
F. Mahmoudi, JHEP **0712**, 026 (2007);  
K.A. Olive and L. Velasco-Sevilla, JHEP **0805**, 052 (2008).
180. S.R. Choudhury and N. Gaur, Phys. Lett. **B451**, 86 (1999);  
K.S. Babu and C.F. Kolda, Phys. Rev. Lett. **84**, 228 (2000);  
G. Isidori and A. Retico, JHEP **0111**, 001 (2001); JHEP **0209**, 063 (2002).
181. The CMS and LHCb Collab., Nature **522**, 68 (2015).
182. K. Hara *et al.* [Belle Collab.], Phys. Rev. Lett. **110**, 131801 (2013);  
J.P. Lees *et al.* [BaBar Collab.], Phys. Rev. **D88**, 031102 (R)(2013).
183. J.P. Lees *et al.* [BaBar Collab.], Phys. Rev. Lett. **109**, 101802 (2012);  
Phys. Rev. **D88**, 072012 (2013).
184. R. Aaij *et al.* [LHCb Collab.], Phys. Rev. Lett. **115**, 111803 (2015);  
M. Huschle *et al.* [Belle Collab.], Phys. Rev. **D92**, 072014 (2015).
185. F. Mahmoudi, S. Neshatpour, and J. Orloff, JHEP **1208**, 092 (2012);  
A. Arbey *et al.*, Phys. Rev. **D87**, 035026 (2013).
186. J. Erler and A. Freitas, “Electroweak Model and Constraints on New Physics,” in the web edition of the *Review of Particle Physics* at <http://pdg.lbl.gov>.
187. J.R. Ellis *et al.*, JHEP **0708**, 083 (2007);  
S. Heinemeyer *et al.*, JHEP **0808**, 087 (2008);  
G.-C. Cho *et al.*, JHEP **1111**, 068 (2011).
188. See the section on neutrinos in “Particle Listings—Leptons” in the web edition of the *Review of Particle Physics* at <http://pdg.lbl.gov>;  
An updated global analysis of neutrino oscillation measurements is provided by the NuFIT collaboration at <http://www.nu-fit.org>.
189. F. Capozzi *et al.*, Phys. Rev. **D89**, 093018 (2014);  
D.V. Forero, M. Tortola and J.W.F. Valle, Phys. Rev. **D90**, 093006 (2014);  
J. Bergstrom *et al.*, JHEP **1509**, 200 (2015).
190. K. Zuber, Phys. Reports **305**, 295 (1998);  
S.F. King, J. Phys. **G42**, 123001 (2015).
191. For a review of neutrino masses in supersymmetry, see *e.g.*, B. Mukhopadhyaya, *Proc. Indian National Science Academy* **A70**, 239 (2004);  
M. Hirsch and J.W.F. Valle, New J. Phys. **6**, 76 (2004).
192. F. Borzumati and Y. Nomura, Phys. Rev. **D64**, 053005 (2001).
193. P. Minkowski, Phys. Lett. **67B**, 421 (1977);  
M. Gell-Mann, P. Ramond, and R. Slansky, in *Supergravity*, edited by D. Freedman and P. van Nieuwenhuizen (North Holland, Amsterdam, 1979) p. 315;  
T. Yanagida, Prog. Theor. Phys. **64**, 1103 (1980);  
R. Mohapatra and G. Senjanovic, Phys. Rev. Lett. **44**, 912 (1980); Phys. Rev. **D23**, 165 (1981).
194. J. Hisano *et al.*, Phys. Lett. **B357**, 579 (1995);  
J. Hisano *et al.*, Phys. Rev. **D53**, 2442 (1996);  
J.A. Casas and A. Ibarra, Nucl. Phys. **B618**, 171 (2001);  
J. Ellis *et al.*, Phys. Rev. **D66**, 115013 (2002);  
A. Masiero, S.K. Vempati, and O. Vives, New J. Phys. **6**, 202 (2004);  
E. Arganda *et al.*, Phys. Rev. **D71**, 035011 (2005);  
F.R. Joaquim and A. Rossi, Phys. Rev. Lett. **97**, 181801 (2006);  
J.R. Ellis and O. Lebedev, Phys. Lett. **B653**, 411 (2007).
195. Y. Grossman and H.E. Haber, Phys. Rev. Lett. **78**, 3438 (1997);  
A. Dedes, H.E. Haber, and J. Rosiek, JHEP **0711**, 059 (2007).
196. M. Hirsch, H.V. Klapdor-Kleingrothaus, and S.G. Kovalenko, Phys. Lett. **B398**, 311 (1997);  
L.J. Hall, T. Moroi, and H. Murayama, Phys. Lett. **B424**, 305 (1998);  
K. Choi, K. Hwang, and W.Y. Song, Phys. Rev. Lett. **88**, 141801 (2002);  
T. Honkavaara, K. Huitu, and S. Roy, Phys. Rev. **D73**, 055011 (2006).
197. M. Chemtob, Prog. in Part. Nucl. Phys. **54**, 71 (2005);  
R. Barbier *et al.*, Phys. Reports **420**, 1 (2005).
198. H. Dreiner, in *Perspectives on Supersymmetry II*, edited by G.L. Kane (World Scientific, Singapore, 2010) pp. 565–583.
199. B.C. Allanach, A. Dedes, and H.K. Dreiner, Phys. Rev. **D60**, 075014 (1999).
200. L.E. Ibáñez and G.G. Ross, Nucl. Phys. **B368**, 3 (1992);  
L.E. Ibáñez, Nucl. Phys. **B398**, 301 (1993).
201. A. Dedes, S. Rimmer, and J. Rosiek, JHEP **0608**, 005 (2006);  
B.C. Allanach and C.H. Kom, JHEP **0804**, 081 (2008);  
H.K. Dreiner *et al.*, Phys. Rev. **D84**, 113005 (2011).
202. L.E. Ibáñez and G.G. Ross, Nucl. Phys. **B368**, 3 (1992).
203. H.K. Dreiner, C. Luhn, and M. Thormeier Phys. Rev. **D73**, 075007 (2006).
204. K. Tamvakis, Phys. Lett. **B382**, 251 (1996);  
G. Eyal and Y. Nir, JHEP **9906**, 024 (1999);  
A. Florex *et al.*, Phys. Rev. **D87**, 095010 (2013).
205. B.C. Allanach and B. Gripaios, JHEP **1205**, 062 (2012);  
M. Asano, K. Rolbiecki, and K. Sakurai, JHEP **1301**, 128 (2013).
206. See *e.g.*, J.C. Romao, Nucl. Phys. (Proc. Supp.) **81**, 231 (2000);  
Y. Grossman and S. Rakshit, Phys. Rev. **D69**, 093002 (2004).
207. R.N. Mohapatra, Phys. Rev. **D34**, 3457 (1986);  
K.S. Babu and R.N. Mohapatra, Phys. Rev. Lett. **75**, 2276 (1995);  
M. Hirsch, H.V. Klapdor-Kleingrothaus, and S.G. Kovalenko, Phys. Rev. Lett. **75**, 17 (1995); Phys. Rev. **D53**, 1329 (1996).
208. Y. Grossman and H.E. Haber, Phys. Rev. **D59**, 093008 (1999).
209. S. Dimopoulos and L.J. Hall, Phys. Lett. **B207**, 210 (1988);  
J. Kalinowski *et al.*, Phys. Lett. **B406**, 314 (1997);  
J. Erler, J.L. Feng, and N. Polonsky, Phys. Rev. Lett. **78**, 3063 (1997).
210. H.K. Dreiner, P. Richardson, and M.H. Seymour, Phys. Rev. **D63**, 055008 (2001).
211. J.E. Kim and H.P. Nilles, Phys. Lett. **B138**, 150 (1984).
212. J.E. Kim and H.P. Nilles, Mod. Phys. Lett. **A9**, 3575 (1994).

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## Searches Particle Listings

### Supersymmetric Particle Searches

213. G.F. Giudice and A. Masiero, Phys. Lett. **B206**, 480 (1988).
214. J.A. Casas and C. Munoz, Phys. Lett. **B306**, 288 (1993).
215. M. Cvetič *et al.*, Phys. Rev. **D56**, 2861 (1997) [Erratum: **D58**, 119905 (1998)].
216. For a review see e.g., R. D. Peccei, Lect. Notes Phys. **741**, 3 (2008).
217. R. Peccei and H. Quinn, Phys. Rev. Lett. **38**, 1440 (1977); Phys. Rev. **D16**, 1791 (1977).
218. H. Murayama, H. Suzuki, and T. Yanagida, Phys. Lett. **B291**, 418 (1992); T. Gherghetta, G.L. Kane, Phys. Lett. **B354**, 300 (1995); K.J. Bae, H. Baer, and H. Serce, Phys. Rev. **D91**, 015003 (2015).
219. A. Delgado, G. Nardini, and M. Quiros, Phys. Rev. **D86**, 115010 (2012).
220. P. Fayet, Phys. Lett. **78B**, 417 (1978).
221. For a review, see *e.g.*, K. Benakli, Fortsch. Phys. **59**, 1079 (2011).
222. P.J. Fox, A.E. Nelson, and N. Weiner, JHEP **0208**, 035 (2002).
223. K. Benakli and M.D. Goodsell, Nucl. Phys. **B816**, 185 (2009); Nucl. Phys. **B840**, 1 (2010).
224. G.D. Kribs, E. Poppitz, and N. Weiner, Phys. Rev. **D78**, 055010 (2008).
225. K. Benakli, M.D. Goodsell, and F. Staub, JHEP **1306**, 073 (2013).
226. See *e.g.*, J.L. Hewett and T.G. Rizzo, Phys. Reports **183**, 193 (1989).
227. S.F. King, S. Moretti, and R. Nevzorov, Phys. Lett. **B634**, 278 (2006); Phys. Rev. **D73**, 035009 (2006).

#### SUPERSYMMETRY, PART II (EXPERIMENT)

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#### II.1. Introduction

Supersymmetry (SUSY), a transformation relating fermions to bosons and vice versa [1–9], is one of the most compelling possible extensions of the Standard Model of particle physics (SM) that could be discovered at high-energy colliders such as the Large Hadron Collider (LHC) at CERN.

On theoretical grounds SUSY is motivated as a generalization of space-time symmetries. A low-energy realization of SUSY, *i.e.*, SUSY at the TeV scale, is, however, not a necessary consequence. Instead, low-energy SUSY is motivated by the possible cancellation of quadratic divergences in radiative corrections to the Higgs boson mass [10–15]. Furthermore, it is intriguing that a weakly interacting, (meta)stable supersymmetric particle might make up some or all of the dark matter in the universe [16–18]. In addition, SUSY predicts that gauge couplings, as measured experimentally at the electroweak scale, unify at an energy scale  $\mathcal{O}(10^{16})\text{GeV}$  (“GUT scale”) near the Planck scale [19–25].

In the minimal supersymmetric extension to the Standard Model, the so called MSSM [11,26,27], a supersymmetry transformation relates every fermion and gauge boson in the SM to a supersymmetric partner with half a unit of spin difference, but otherwise with the same properties and quantum numbers. These are the “sfermions”: squarks ( $\tilde{q}$ ) and sleptons ( $\tilde{\ell}$ ,  $\tilde{\nu}$ ), and the “gauginos”. The MSSM Higgs sector contains two doublets, for up-type quarks and for down-type quarks and charged leptons respectively. After electroweak symmetry breaking, five Higgs bosons arise, of which two are charged. The supersymmetric partners of the Higgs doublets are known as “higgsinos.” The charged weak gauginos and higgsinos mix to “charginos” ( $\tilde{\chi}^\pm$ ), and the neutral ones mix to “neutralinos” ( $\tilde{\chi}^0$ ). The SUSY partners of the gluons are known as “gluinos” ( $\tilde{g}$ ). The fact that such particles are not yet observed leads to the conclusion that, if supersymmetry is realized, it is a broken symmetry. A description of SUSY in the form of an effective Lagrangian with only “soft” SUSY breaking terms and SUSY masses at the TeV scale maintains cancellation of quadratic divergences in particle physics models.

The phenomenology of SUSY is to a large extent determined by the SUSY breaking mechanism and the SUSY breaking scale. This determines the SUSY particle masses, the mass hierarchy, the field contents of physical particles, and their decay modes. In addition, phenomenology crucially depends on whether the multiplicative quantum number of R-parity [27],  $R = (-1)^{3(B-L)+2S}$ , where  $B$  and  $L$  are baryon and lepton numbers and  $S$  is the spin, is conserved or violated. If R-parity is conserved, SUSY particles (sparticles), which have odd R-parity, are produced in pairs and the decays of each SUSY particle must involve an odd number of lighter SUSY particles. The lightest SUSY particle (LSP) is then stable and often assumed to be a weakly interacting massive particle (WIMP). If R-parity is violated, new terms  $\lambda_{ijk}$ ,  $\lambda'_{ijk}$  and  $\lambda''_{ijk}$  appear in the superpotential, where  $ijk$  are generation indices;  $\lambda$ -type couplings appear between lepton superfields

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only,  $\lambda''$ -type are between quark superfields only, and  $\lambda'$ -type couplings connect the two. R-parity violation implies lepton and/or baryon number violation. More details of the theoretical framework of SUSY are discussed elsewhere in this volume [28].

Today low-energy data from flavor physics experiments, high-precision electroweak observables as well as astrophysical data impose strong constraints on the allowed SUSY parameter space. Recent examples of such data include measurements of the rare B-meson decay  $B_s \rightarrow \mu^+ \mu^-$  [29] and accurate determinations of the cosmological dark matter relic density constraint [30,31].

These indirect constraints are often more sensitive to higher SUSY mass scales than experiments searching for direct sparticle production at colliders, but the interpretation of these results is often strongly model dependent. In contrast, direct searches for sparticle production at collider experiments are less subject to interpretation ambiguities and therefore they play a crucial role in the search for SUSY.

The discovery of a new scalar boson with a mass around 125 GeV compatible with a Higgs boson imposes constraints on SUSY, which are discussed elsewhere [28,32].

In this review we limit ourselves to direct searches, covering data analyses at LEP, HERA, the Tevatron and the LHC, with emphasis on the latter. For more details on LEP and Tevatron constraints, see earlier PDG reviews [33].

### II.2. Experimental search program

The electron-positron collider LEP was operational at CERN between 1989 and 2000. In the initial phase, center-of-mass energies around the  $Z$ -peak were probed, but after 1995 the LEP experiments collected a significant amount of luminosity at higher center-of-mass energies, some  $235 \text{ pb}^{-1}$  per experiment at  $\sqrt{s} \geq 204 \text{ GeV}$ , with a maximum  $\sqrt{s}$  of 209 GeV.

Searches for new physics at  $e^+e^-$  colliders benefit from the clean experimental environment and the fact that momentum balance can be measured not only in the plane transverse to the beam, but also in the direction along the beam (up to the beam pipe holes), defined as the longitudinal direction. Searches at LEP are dominated by the data samples taken at the highest center-of-mass energies.

Significant constraints on SUSY have been set by the CDF and D0 experiments at the Tevatron, a proton-antiproton collider at a center-of-mass energy of up to 1.96 TeV. CDF and D0 have collected integrated luminosities between 10 and  $11 \text{ fb}^{-1}$  each up to the end of collider operations in 2011.

The electron-proton collider HERA provided collisions to the H1 and ZEUS experiments between 1992 and 2007, at a center-of-mass energy up to 318 GeV. A total integrated luminosity of approximately  $0.5 \text{ fb}^{-1}$  has been collected by each experiment. Since in  $ep$  collisions no annihilation process takes place, SUSY searches at HERA typically look for R-parity violating production of single SUSY particles.

The LHC has started proton-proton operation at a center-of-mass energy of 7 TeV in 2010. By the end of 2011 the experiments ATLAS and CMS had collected about  $5 \text{ fb}^{-1}$  of integrated luminosity each, and the LHCb experiment had collected approximately  $1 \text{ fb}^{-1}$ . In 2012, the LHC operated at a center-of-mass energy of 8 TeV, and ATLAS and CMS collected approximately  $20 \text{ fb}^{-1}$  each, whereas LHCb collected  $2 \text{ fb}^{-1}$ . In 2015, the LHC has started Run 2, with center-of-mass energy of 13 TeV.

Proton-(anti)proton colliders produce interactions at higher center-of-mass energies than those available at LEP, and cross sections of QCD-mediated processes are larger, which is reflected in the higher sensitivity for SUSY particles carrying color charge: squarks and gluinos. Large background contributions from Standard Model processes, however, pose challenges to trigger and analysis. Such backgrounds are dominated by multijet production processes, including, particularly at the LHC, those of top quark production, as well as jet production in association with vector bosons. The proton momentum is shared between its parton constituents, and in each collision only a fraction of the total center-of-mass energy is available in the hard parton-parton scattering. Since the parton momenta in the longitudinal direction are not known on an event-by-event basis, use of momentum conservation constraints in an analysis is restricted to the transverse plane, leading to the definition of transverse variables, such as the missing transverse momentum, and the transverse mass. Proton-proton collisions at the LHC differ from proton-antiproton collisions at the Tevatron in the sense that there are no valence anti-quarks in the proton, and that gluon-initiated processes play a more dominant role. The increased center-of-mass energy of the LHC compared to the Tevatron significantly extends the kinematic reach for SUSY searches. This is reflected foremost in the sensitivity for squarks and gluinos, but also for other SUSY particles.

The main production mechanisms of massive colored sparticles at hadron colliders are squark-squark, squark-gluino and gluino-gluino production; when “squark” is used “antisquark” is also implied. The typical SUSY search signature at hadron colliders contains high- $p_T$  jets, which are produced in the decay chains of heavy squarks and gluinos, and significant missing momentum originating from the two LSPs produced at the end of the decay chain. Assuming R-parity conservation, the LSPs are neutral and weakly interacting massive particles which escape detection. Standard Model backgrounds with missing transverse momentum include leptonic  $W/Z$ -boson decays, heavy-flavor decays to neutrinos, and multijet events that may be affected by instrumental effects such as jet mismeasurement.

Selection variables designed to separate the SUSY signal from the Standard Model backgrounds include  $H_T$ ,  $E_T^{\text{miss}}$ , and  $m_{\text{eff}}$ . The quantities  $H_T$  and  $E_T^{\text{miss}}$  refer to the measured transverse energy and missing transverse momentum in the event, respectively. They are usually defined as the scalar ( $H_T$ ) and negative vector sum ( $E_T^{\text{miss}}$ ) of the transverse jet energies or transverse calorimeter clusters energies measured in



the event. The quantity  $m_{\text{eff}}$  is referred to as the effective mass of the event and is defined as  $m_{\text{eff}} = H_T + |E_T^{\text{miss}}|$ . The peak of the  $m_{\text{eff}}$  distribution for SUSY signal events correlates with the SUSY mass scale, in particular with the mass difference between the primary produced SUSY particle and the LSP [34], whereas the Standard Model backgrounds dominate at low  $m_{\text{eff}}$ . Additional reduction of multijet backgrounds can be achieved by demanding isolated leptons or photons in the final states; in such events the lepton or photon transverse momentum may be added to  $H_T$  or  $m_{\text{eff}}$  for further signal-background separation.

In the past few years alternative approaches have been developed to increase the sensitivity to pair production of heavy sparticles with masses around 1 TeV focusing on the kinematics of their decays, and to further suppress the background from multijet production. Prominent examples of these new approaches are searches using the  $\alpha_T$  [35–39], *razor* [40], *stransverse mass* ( $m_{T2}$ ) [41], and *contransverse mass* ( $m_{CT}$ ) [42] variables.

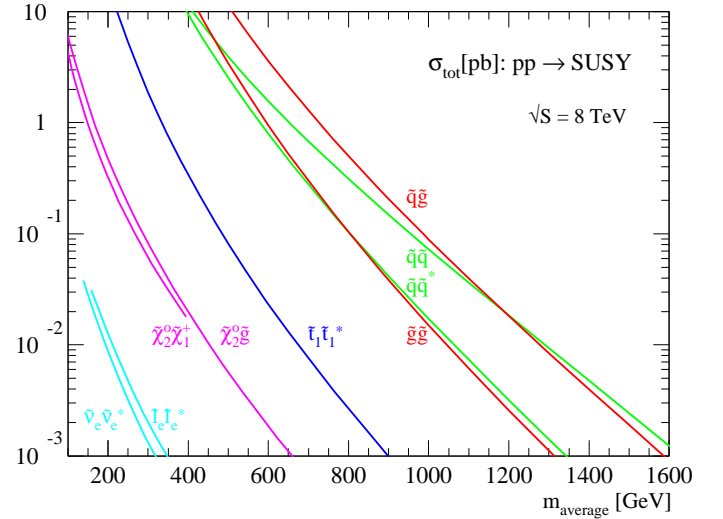
### II.3. Interpretation of results

Since the mechanism by which SUSY is broken is unknown, a general approach to SUSY via the most general soft SUSY breaking Lagrangian adds a significant number of new free parameters. For the minimal supersymmetric standard model, MSSM, *i.e.*, the model with the minimal particle content, these comprise 105 new parameters. A phenomenological analysis of SUSY searches leaving all these parameters free is not feasible. For the practical interpretation of SUSY searches at colliders several approaches are taken to reduce the number of free parameters.

One approach is to assume a SUSY breaking mechanism and lower the number of free parameters through the assumption of additional constraints. In particular in past years, interpretations of experimental results were predominately performed in constrained models of gravity mediated [43,44], gauge mediated [45,46], and anomaly mediated [47,48] SUSY breaking. Before the start of the LHC and even during its first year of operation, the most popular model for interpretation of collider based SUSY searches was the constrained MSSM (CMSSM) [43,49,50], which in the literature is also referred to as minimal supergravity, or MSUGRA. The CMSSM is described by five parameters: the common sfermion mass  $m_0$ , the common gaugino mass  $m_{1/2}$ , and the common trilinear coupling parameter  $A_0$ , all defined at the GUT scale, the ratio of the vacuum expectation values of the Higgs fields for up-type and down-type fermions  $\tan\beta$ , and the sign of the higgsino mass parameter  $\mu$ , defined at the electroweak scale. In gauge mediation models, the paradigm of general gauge mediation (GGM) [51] is slowly replacing minimal gauge mediation, denoted traditionally as GMSB (gauge mediated SUSY breaking).

These constrained SUSY models are theoretically well motivated and provide a rich spectrum of experimental signatures. Therefore, they represent a useful framework to benchmark performance, compare limits or reaches and assess the expected

sensitivity of different search strategies. However, with universality relations imposed on the soft SUSY breaking parameters, they do not cover all possible kinematic signatures and mass relations of SUSY. In such scenarios the squarks are often nearly degenerate in mass, in particular for the first and second generation. The exclusion of parameter space in the CMSSM and in CMSSM-inspired models is mainly driven by first and second generation squark production together with gluino production.



**Figure 1:** Cross sections for pair production of different sparticles as a function of their mass at the LHC for a center-of-mass energy of 8 TeV [52]. Typically the production cross section of colored squarks and gluinos is several orders of magnitude larger than the one for leptons or charginos. Except for the explicitly shown pair production of stops, production cross sections for squarks assumes mass degeneracy of left- and right-handed  $u$ ,  $d$ ,  $s$ ,  $c$  and  $b$  squarks.

As shown in Fig. 1 [52] these processes possess the largest production cross sections in proton-proton collisions, and thus the LHC searches typically provide the tightest mass limits on these colored sparticles. This, however, implies that the allowed parameter space of constrained SUSY models today has been restrained significantly by searches from ATLAS and CMS. Furthermore, confronting the remaining allowed parameter space with other collider and non-collider measurements, which are directly or indirectly sensitive to contributions from SUSY, the overall compatibility of these models with all data is significantly worse than in the pre-LHC era (see section II.7 for further discussion), indicating that very constrained models like the CMSSM might no longer be good benchmark scenarios to solely characterize the results of SUSY searches at the LHC.

For these reasons, an effort has been made in the past years to complement the traditional constrained models with more flexible approaches.

One answer to study a broader and more comprehensive subset of the MSSM is via the phenomenological-MSSM, or

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pMSSM [53–55]. It is derived from the MSSM, using experimental data to eliminate parameters that are free in principle but have already been highly constrained by measurements of *e.g.*, flavor mixing and CP-violation. This effective approach reduces the number of free parameters in the MSSM to typically 19, making it a practical compromise between the full MSSM and highly constrained models such as the CMSSM.

Even less dependent on fundamental assumptions are interpretations in terms of so-called simplified models [56–59]. Such models assume a limited set of SUSY particle production and decay modes and leave open the possibility to vary masses and other parameters freely. Therefore, simplified models enable comprehensive studies of individual SUSY topologies, and are useful for optimization of the experimental searches over a wide parameter space. As a consequence, since 2011 ATLAS and CMS have adopted simplified models as the primary framework to provide interpretations of their searches. Today, almost every individual search provides interpretations of their results in one or even several simplified models that are characteristic of SUSY topologies probed by the analysis.

However, while these models are very convenient for the interpretation of individual SUSY production and decay topologies, care must be taken when applying these limits to more complex SUSY spectra. Therefore, in practise, simplified model limits are often used as an approximation of the constraints that can be placed on sparticle masses in more complex SUSY spectra. Yet, depending on the assumed SUSY spectrum, the sparticle of interest, and the considered simplified model limit, this approximation can lead to a significant mistake, typically an overestimation, in the assumed constraint on the sparticle mass (see for example [60]). Only on a case-by-case basis can it be determined whether the limit of a given simplified model represents a good approximation of the true underlying constraint that can be applied on a sparticle mass in a complex SUSY spectrum. In the following, we will always point out explicitly the assumptions that have entered the limits when quoting interpretations from simplified models.

This review covers results up to September 2015 and since none of the searches performed so far have shown significant excess above the SM background prediction, the interpretation of the presented results are exclusion limits on SUSY parameter space.

### II.4. Exclusion limits on gluino and squark masses

Gluinos and squarks are the SUSY partners of gluons and quarks, and thus carry color charge. Limits on squark masses of the order 100 GeV have been set by the LEP experiments [61]. However, due to the colored production of these particles at hadron colliders (see *e.g.* Fig. 1), hadron collider experiments are able to set much tighter mass limits.

Pair production of these massive colored sparticles at hadron colliders generally involve both s-channel and t-channel parton-parton interactions. Since there is a negligible amount of bottom and top quark content in the proton, top- and bottom

squark production proceeds through s-channel diagrams only with smaller cross sections. In the past, experimental analyses of squark and/or gluino production typically assumed the first and second generation squarks to be approximately degenerate in mass. However, in order to have even less model dependent interpretations of the searches, the experiments have started to also provide simplified model limits on individual first or second generation squarks.

Assuming R-parity conservation and assuming gluinos to be heavier than squarks, squarks will predominantly decay to a quark and a neutralino or chargino, if kinematically allowed. The decay may involve the lightest neutralino (typically the LSP) or chargino, but, depending on the masses of the gauginos, may involve heavier neutralinos or charginos. For pair production of first and second generation squarks, the simplest decay modes involve two jets and missing momentum, with potential extra jets stemming from initial state or final state radiation (ISR/FSR) or from decay modes with longer cascades. Similarly, gluino pair production leads to four jets and missing momentum, and possibly additional jets from ISR/FSR or cascades. Associated production of a gluino and a (anti-)squark is also possible, in particular if squarks and gluinos have similar masses, typically leading to three or more jets in the final state. In cascades, isolated photons or leptons may appear from the decays of sparticles such as neutralinos or charginos. Final states are thus characterized by significant missing transverse momentum, and at least two, and possibly many more high  $p_T$  jets, which can be accompanied by one or more isolated objects like photons or leptons, including  $\tau$  leptons, in the final state. Table 1 shows a schematic overview of characteristic final state signatures of gluino and squark production for different mass hierarchy hypotheses and assuming decays involving the lightest neutralino.

The main reference for ATLAS results in this section is the ATLAS Run 1 summary paper on squark and gluino production [62], while for CMS individual results are cited.

**Table 1:** Typical search signatures at hadron colliders for direct gluino and first- and second-generation squark production assuming different mass hierarchies.

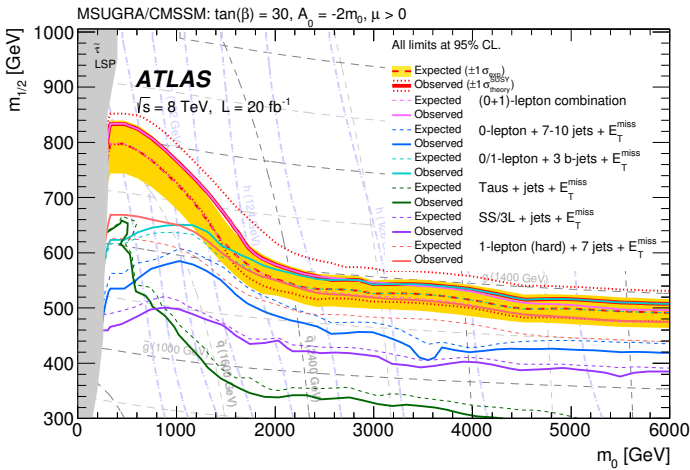
Mass Hierarchy	Main Production	Dominant Decay	Typical Signature
$m_{\tilde{q}} \ll m_{\tilde{g}}$	$\tilde{q}\tilde{q}, \tilde{q}\tilde{\bar{q}}$	$\tilde{q} \rightarrow q\tilde{\chi}_1^0$	$\geq 2 \text{ jets} + E_T^{\text{miss}} + X$
$m_{\tilde{q}} \approx m_{\tilde{g}}$	$\tilde{q}\tilde{g}, \tilde{q}\tilde{\bar{g}}$	$\tilde{q} \rightarrow q\tilde{\chi}_1^0$ $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	$\geq 3 \text{ jets} + E_T^{\text{miss}} + X$
$m_{\tilde{q}} \gg m_{\tilde{g}}$	$\tilde{g}\tilde{g}$	$\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	$\geq 4 \text{ jets} + E_T^{\text{miss}} + X$

#### II.4.1 Exclusion limits on the gluino mass

Limits set by the Tevatron experiments on the gluino mass assume the framework of the CMSSM, with  $\tan\beta = 5$  (CDF) or  $\tan\beta = 3$  (D0),  $A_0 = 0$  and  $\mu < 0$ , and amount to lower limits

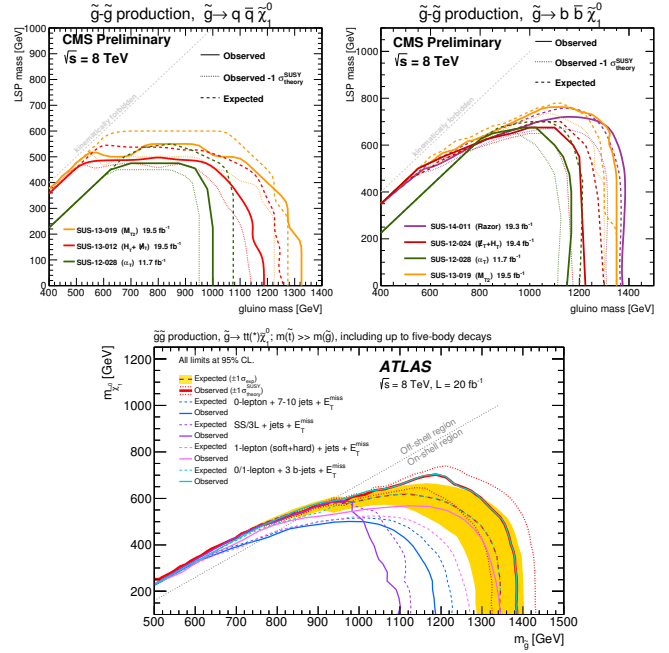
of about 310 GeV for all squark masses, or 390 GeV for the case  $m_{\tilde{q}} = m_{\tilde{g}}$  [63,64]. During the first year of physics operation of the LHC in 2010, these limits have been superseded by those provided by ATLAS and CMS.

Today, limits on the gluino mass have been set using up to approximately  $20 \text{ fb}^{-1}$  of data recorded at a center-of-mass energy of 8 TeV. As shown in Fig. 2, the ATLAS collaboration places limits for several searches in the framework of the CMSSM, assuming  $\tan\beta = 30$ ,  $A_0 = -2m_0$ , and  $\mu > 0$  [62]; these parameter values are chosen since they lead to a mass for the lightest Higgs boson compatible with 125 GeV in a large part of the  $m_0 - m_{1/2}$  plane. For low  $m_0$  the combination of the 0+1 lepton plus jets and missing momentum analyses provides the best sensitivity, while for values of  $m_0$  above  $\approx 1800$  GeV a more dedicated search requiring zero or one isolated lepton accompanied with at least three jets identified to originate from bottom quarks ( $b$ -jets) takes over. The limits at low  $m_0$  are mainly driven by squark-gluino and squark-squark production and at high  $m_0$  gluino pair production dominates. As also indicated in Fig. 1, all other particle production modes do not play a significant role for limits in the CMSSM. In this constrained model gluino masses below around 1300 GeV are excluded by the ATLAS collaboration for all squark masses, while for equal squark and gluino masses, the limit is about 1700 GeV. Further details about the searches that are displayed in Fig. 2 can be obtained from the ATLAS summary paper on squark and gluino production [62].



**Figure 2:** Limits, at 95% C.L., derived from several different ATLAS searches in the CMSSM parameters  $m_0$  and  $m_{1/2}$ , assuming  $\tan\beta = 30$ ,  $A_0 = -2m_0$  and  $\mu > 0$  [62].

Limits on the gluino mass have also been established in the framework of simplified models. Assuming only gluino pair production, in particular three primary decay chains of the gluino have been considered by the LHC experiments for interpretations of their search results. The first decay chain



**Figure 3:** Upper mass limits, at 95% C.L., on gluino pair production for the decay chains  $\tilde{g} \rightarrow q\bar{q}\chi_1^0$  (upper left) [65],  $\tilde{g} \rightarrow b\bar{b}\chi_1^0$  (upper right) [66], and  $\tilde{g} \rightarrow t\bar{t}\chi_1^0$  (lower middle) [62,157]. The limits are defined in the framework of simplified models assuming a single decay chain, (i.e. 100% branching fraction). The upper plots show limits from the CMS collaboration, while the displayed limits for  $\tilde{g} \rightarrow t\bar{t}\chi_1^0$  are obtained from ATLAS searches. It should be noted that the ATLAS results include both on-shell as well as off-shell production of gluino induced stop production (see text for details).

$\tilde{g} \rightarrow q\bar{q}\chi_1^0$  assumes gluino mediated production of first and second generation squarks which leads to four light flavor quarks in the final state. Therefore, inclusive all-hadronic analyses searching for multijet plus  $E_T^{\text{miss}}$  final states are utilized to put limits on this simplified model. These limits are derived as a function of the gluino and neutralino (LSP) mass. As shown in Fig. 3 (upper left), using the cross section from next-to-leading order QCD corrections and the resummation of soft gluon emission at next-to-leading-logarithmic accuracy as reference, the CMS collaboration [65] excludes in this simplified model gluino masses below approximately 1300 GeV (see also [66–68]), for a massless neutralino. In scenarios where neutralinos are not very light, the efficiency of the analyses is reduced by the fact that jets are less energetic, and there is less missing transverse momentum in the event. This leads to weaker limits when the mass difference  $\Delta m = m_{\tilde{g}} - m_{\chi_1^0}$  is reduced. For example, for neutralino masses above about 550 GeV no limit on the gluino mass can be set for this decay chain. Therefore, limits on gluino masses are strongly affected by the assumption of the neutralino mass. Similar results for this simplified model have been obtained by ATLAS [62].

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The second important decay chain of the gluino considered for interpretation in a simplified model is  $\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$ . Here the decay is mediated via bottom squarks and thus leads to four jets from  $b$  quarks and  $E_T^{\text{miss}}$  in the final state. Also for this topology inclusive all-hadronic searches provide the highest sensitivity. However, with four  $b$  quarks in the final state, the use of secondary vertex reconstruction for the identification of jets originating from  $b$  quarks provides a powerful handle on the SM background. Therefore, in addition to a multijet plus  $E_T^{\text{miss}}$  signature these searches also require several jets to be tagged as  $b$ -jets. As shown in Fig. 3 (upper right), for this simplified model CMS [66] excludes gluino masses below  $\approx 1350$  GeV for a massless neutralino, while for neutralino masses above  $\approx 750$  GeV no limit on the gluino mass can be set. Comparable limits for this simplified model are provided by searches from ATLAS [62].

Not only first and second generation squarks or bottom squarks may be the product of gluino decays but also, if kinematically allowed, top squarks via the decay  $\tilde{g} \rightarrow \tilde{t}\bar{t}$ . This leads to a “four tops” final state  $ttt\tilde{\chi}_1^0\tilde{\chi}_1^0$  and defines the third important simplified model,  $\tilde{g} \rightarrow \tilde{t}\bar{t}\tilde{\chi}_1^0$ , characterizing gluino pair production. The topology of this decay is very rich in different experimental signatures: as many as four isolated leptons, four  $b$ -jets, several light flavor quark jets, and significant missing momentum from the neutrinos in the  $W$  decay and from the two neutralinos. Therefore, in contrast to the other two simplified models, dedicated searches optimized for this particular final state provide the best mass limit on the gluino for this simplified model. As shown in Fig. 3 (lower middle) [62,157], the ATLAS search requiring significant  $E_T^{\text{miss}}$ , zero or one isolated lepton, and at least three jets identified as  $b$ -jets [69] provides the strongest limit on the gluino mass in the on-shell region ( $m_{\tilde{g}} > 2m_t + m_{\tilde{\chi}_1^0}$ ). At 95% C.L. it rules out a gluino mass below  $\approx 1400$  GeV for  $m_{\tilde{\chi}_1^0} < 300$  GeV. For neutralino masses above  $\approx 700$  GeV, no limit can be placed on the gluino mass for this simplified model. A CMS search [70] also especially optimized for this decay topology by requiring one isolated lepton and high jet multiplicity obtains similar limits.

The ATLAS collaboration also provides limits for the off-shell region ( $m_{\tilde{g}} < 2m_t + m_{\tilde{\chi}_1^0}$ ) of this decay. In the regions of parameter space where the mass difference between the gluino and the lightest neutralino is small, additional sensitivity is obtained from the same-sign search requiring 3 leptons in the final state. To place limits in this off-shell region only four-body ( $\tilde{g} \rightarrow tWb\tilde{\chi}_1^0$ ) and five-body ( $\tilde{g} \rightarrow WbWb\tilde{\chi}_1^0$ ) are considered as for higher multiplicities the gluinos do not decay promptly anymore and thus lead to a different signature topology. With this approach additional parameter space for gluino masses below about 950 GeV can be excluded in the  $m_{\tilde{g}} < 2m_t + m_{\tilde{\chi}_1^0}$  region.

When comparing the limits in Fig. 3 for the three different simplified models it becomes apparent that more parameter space can be excluded when the gluino decay chain is mediated

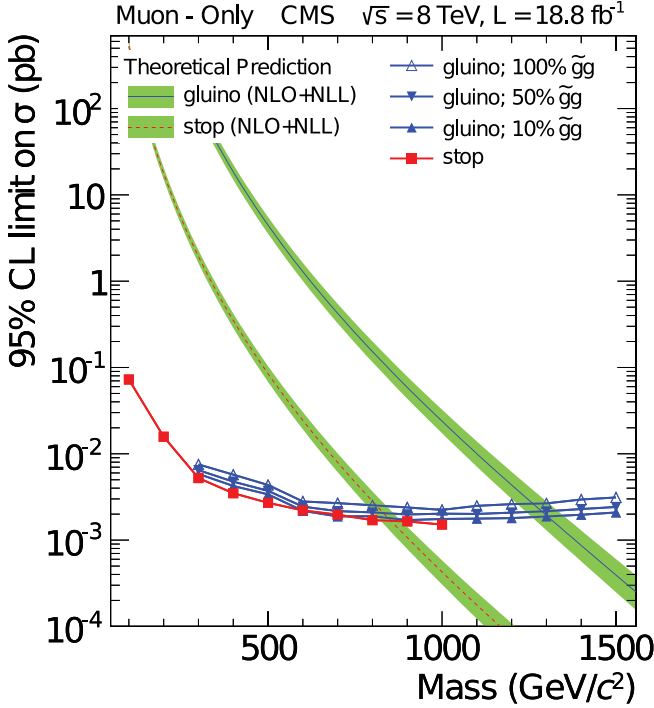
via third generation squarks. The reason for this is the better control of the SM background by means of identification of  $b$ -jets as well as dedicated topology requirements like high jet multiplicity or isolated leptons for these special signatures. However, this variation in sensitivity of the searches for different gluino decay chains is also a clear indication that care must be taken when limits from these simplified models are applied to SUSY models possessing more complex underlying spectra.

If the gluino decay is suppressed, for example if squark masses are high, gluinos may live longer than typical hadronization times. It is expected that such gluinos will hadronize to semi-stable strongly interacting particles known as R-hadrons. Searches for R-hadrons exploit the typical signature of stable charged massive particles in the detector. As shown in Fig. 4, the CMS experiment excludes semi-stable gluino R-hadrons with masses below approximately 1.3 TeV [71]. The limits depend on the probability for gluinos to form bound states known as gluinoballs, as these are neutral and not observed in the tracking detectors. Similar limits are obtained by the ATLAS experiment [72]. Limits of about 1 TeV are set in the scenario of R-hadron decays inside the detector, for  $c\tau$  ranging from 1 to 1000 mm [73].

Alternatively, since such R-hadrons are strongly interacting, they may be stopped in the calorimeter or in other material, and decay later into energetic jets. These decays are searched for by identifying the jets outside the time window associated with bunch-bunch collisions [74–76]. The latest ATLAS analysis [75] based on the full 2011 and 2012 data set combined (28 fb<sup>-1</sup>) places limits at 95% C.L. on gluino production over almost 16 orders of magnitude in gluino lifetime. For  $m_{\tilde{\chi}_1^0} > 100$  GeV, assuming a 100% branching fraction for gluino decay to gluon (or  $q\bar{q}$ ) + neutralino, gluinos with lifetimes from 10  $\mu$ s to 1000 s and  $m_{\tilde{g}} < 832$  GeV are excluded. When SUSY spectra are compressed, this limit weakens to an exclusion of  $m_{\tilde{g}} < 545$  GeV for  $m_{\tilde{g}} - m_{\tilde{\chi}_1^0} < 100$  GeV.

In summary, for interpretations in the CMSSM, simplified models, and semi-stable R-hadrons, the best limits on the gluino mass range from around 1200 GeV to about 1400 GeV, while for interpretations in the context of stopped R-hadrons the limit on  $m_{\tilde{g}}$  is around 850 GeV. All these limits weaken significantly for compressed SUSY spectra when the mass difference  $m_{\tilde{g}} - m_{\tilde{\chi}_1^0}$  is reduced.

R-parity violating gluino decays are searched for in a number of final states. Searches in multilepton final states set lower mass limits of 1 to 1.4 TeV, depending on neutralino mass and lepton flavor, on decays mediated by  $\lambda$  and  $\lambda'$  couplings [77,78], assuming prompt decays. Searches for displaced vertices are sensitive to non-prompt decays [73]. Multijet final states have been used to search for fully hadronic gluino decays involving  $\lambda''$ , by CDF [79], ATLAS [80] and CMS [81]. Lower mass limits range between 800 and 1000 GeV depending on neutralino mass and flavor content of the final state.



**Figure 4:** Observed 95% C.L. upper limits on the cross section for (semi-)stable top squarks or gluinos [71]. For gluinos, different fractions of gluinoball states produced after hadronization scenarios are indicated. The observed limits are compared with the predicted theoretical cross sections where the bands represent the theoretical uncertainties on the cross section values.

#### II.4.2. Exclusion limits on first and second generation squark masses

Limits on first and second generation squark masses set by the Tevatron experiments assume the CMSSM model, and amount to lower limits of about 380 GeV for all gluino masses, or 390 GeV for the case  $m_{\tilde{q}} = m_{\tilde{g}}$  [63,64].

At the LHC, limits on squark masses have been set using up to approximately  $20 \text{ fb}^{-1}$  of data at 8 TeV. As shown in Fig. 2, the ATLAS collaboration [62] excludes in the framework of the CMSSM squark masses below  $\approx 1600$  GeV for all gluino masses. For equal squark and gluino masses, the limit is about 1700 GeV.

Interpretations in simplified models are typically characterizing squark pair production with only one decay chain of  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ . Here it is assumed that the left and right-handed  $\tilde{u}$ ,  $\tilde{d}$ ,  $\tilde{s}$  and  $\tilde{c}$  squarks are degenerate in mass. Furthermore, it is assumed that the mass of the gluino is very high and thus contributions of the corresponding t-channel diagrams to squark pair production are negligible. Therefore, the total production cross section for this simplified model is eight times

the production cross section of an individual squark (e.g.  $\tilde{u}_L$ ). The CMS collaboration provides interpretations using different all-hadronic searches for this simplified model. As displayed in Fig. 5, best observed exclusion is obtained from the analysis using the  $m_{T2}$  variable [65], which excludes squark masses just below 925 GeV for a light neutralino. The effects of heavy neutralinos on squark limits are similar to those discussed in the gluino case (see section II.4.1) and only for neutralino masses below  $\approx 350$  GeV squark masses can be excluded. Results from the ATLAS collaboration [62] for this simplified model are similar.

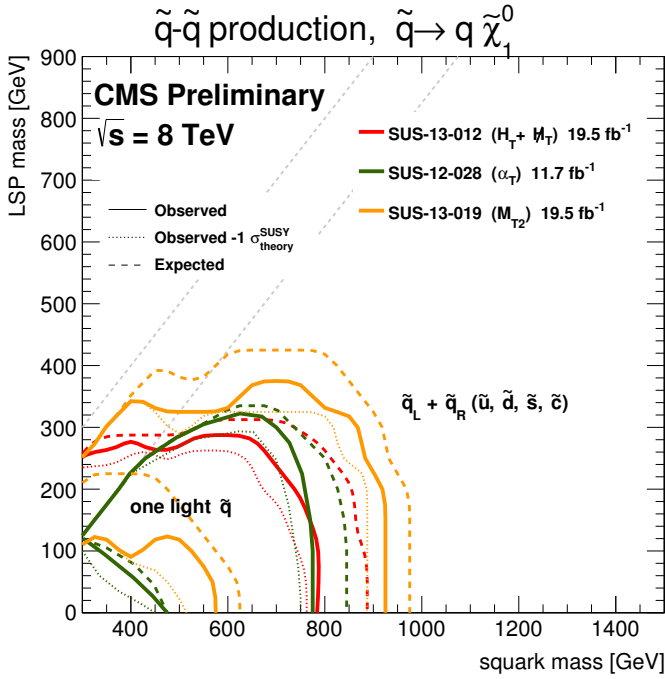
For the same analysis ATLAS also provides an interpretation of their search result in a simplified model assuming strong production of first and second generation squarks in association with gluinos. This interpretation excludes squark masses below  $\approx 1400$  GeV for all gluino masses as well as gluino masses below  $\approx 1400$  GeV for all squark masses. For equal squark and gluino masses, the limit is about 1700 GeV and therefore very similar to limits provided in the CMSSM.

If the assumption of mass degenerate first and second generation squarks is dropped and only the production of a single light squark is assumed, the limits weaken significantly. This is shown as the much smaller exclusion region in Fig. 5, which represents the 95% C.L. upper limit of pair production of a single light squark, with the gluino and all other squarks decoupled to very high masses. With a best observed squark mass limit of only  $\approx 575$  GeV for a massless neutralino and a neutralino mass of  $\approx 120$  GeV above which no squark mass limit can be placed, the exclusion reach of the LHC experiments for single light squark is rather weak. It should be noted that this limit is not a result of a simple scaling of the above mentioned mass limits assuming eightfold mass degeneracy but it also takes into account that for an eight times lower production cross section the analyses must probe kinematic regions of phase space that are closer to the ones of SM background production. Since signal acceptance and the ratio of expected signal to SM background events of the analyses are typically worse in this region of phase space not only the  $1/8$  reduction in production cross section but also a worse analysis sensitivity are responsible for the much weaker limit on single squark pair production.

For single light squarks ATLAS also reports results of a dedicated search for pair production of scalar partners of charm quarks [82]. Assuming that the scalar-charm state exclusively decays into a charm quark and a neutralino, scalar-charm masses up to 490 GeV are excluded for neutralino masses below 200 GeV.

R-parity violating production of single squarks via a  $\lambda'$ -type coupling has been studied at HERA. In such models, a lower limit on the squark mass of the order of 275 GeV has been set for electromagnetic-strength-like couplings  $\lambda' = 0.3$  [83]. At the LHC, both prompt [77,78] and non-prompt [73,84] R-parity





**Figure 5:** 95% C.L. exclusion contours in the squark-neutralino mass plane defined in the framework of simplified models assuming a single decay chain of  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  [65]. Two assumptions for the squark pair production cross sections are displayed; a) eightfold degeneracy for the masses of the first and second generation squarks and b) only one light flavor squark.

violating squark decays have been searched for, but no signal was found. Squark mass limits are very model-dependent.

### II.4.3. Exclusion limits on third generation squark masses

SUSY at the TeV-scale is often motivated by naturalness arguments, most notably as a solution to stabilize quadratic divergences in radiative corrections to the Higgs boson mass. In this context, the most relevant terms for SUSY phenomenology arise from the interplay between the masses of the third generation squarks and the Yukawa coupling of the top quark to the Higgs boson. This motivates a potential constraint on the masses of the top squarks and the left-handed bottom squark. Due to the large top quark mass, significant mixing between  $\tilde{t}_L$  and  $\tilde{t}_R$  is expected, leading to a lighter mass state  $\tilde{t}_1$  and a heavier mass state  $\tilde{t}_2$ . In the MSSM, the lightest top squark ( $\tilde{t}_1$ ) can be the lightest squark.

The discovery of a Higgs boson at a mass around 125 GeV has consequences for third generation squarks in the MSSM, which are discussed elsewhere [28]. As a result, and in the absence of a SUSY discovery so far, searches for third generation squark production have become a major focus of the SUSY search program at the LHC. For this reason direct and gluino

mediated top and/or bottom squark production processes, leading to experimental signatures that are rich in jets originating from bottom quarks, are either subject of re-interpretation of inclusive analyses or targets for dedicated searches. The latter ones have become especially important for searches of direct top squark production.

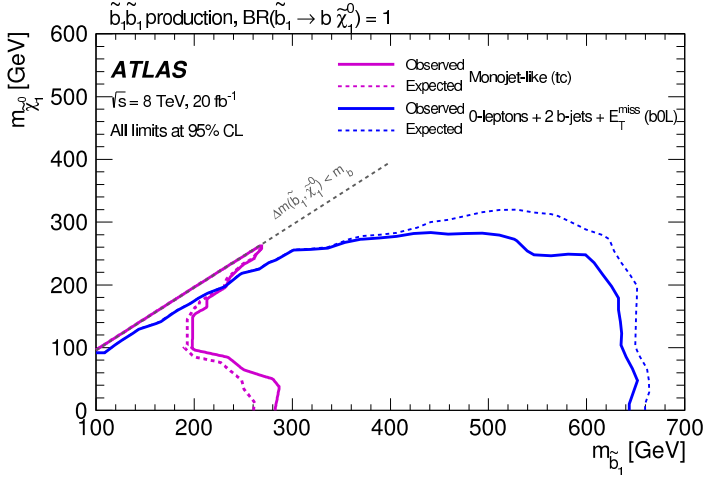
Direct production of top and bottom squark pairs at hadron colliders is suppressed with respect to first generation squarks, due to the absence of  $t$  and  $b$  quarks in the proton (see e.g. the example of direct top squark production in Fig. 1). At the LHC, assuming eightfold mass degeneracy for light flavor squarks as reference, this suppression is at the level of two orders of magnitude for top and bottom squark masses of around 600 GeV. Moreover, at the LHC, there is a very large background of top quark pair production, making especially the experimental analysis of top squark pair production a challenge.

The main reference for ATLAS results in this section is the summary paper for Run1 searches for direct pair production of third-generation squarks [85], while for CMS individual references are cited.

Bottom squarks are expected to decay predominantly to  $b\tilde{\chi}_1^0$  giving rise to the characteristic multi  $b$ -jet and  $E_T^{\text{miss}}$  signature. Direct production of bottom squark pairs has been studied at the Tevatron and at the LHC. Limits from the Tevatron are  $m_{\tilde{b}} > 247$  GeV for a massless neutralino [86,87]. Using the 2011 data the LHC experiments were able to surpass these limits and based on the full 2012 data set, as shown in Fig. 6, using an all-hadronic search requiring significant  $E_T^{\text{miss}}$  and two jets reconstructed as  $b$ -jets in combination with a dedicated mono-jet analysis, ATLAS has set a limit of  $m_{\tilde{b}} \gtrsim 650$  GeV for the same scenario. For  $m_{\tilde{\chi}_1^0} \approx 280$  GeV or higher no limit can be placed on direct bottom squark pair production in this simplified model [85]. The addition of the mono-jet topology increases the sensitivity for compressed spectra allowing for an exclusion of up to  $m_{\tilde{b}} \approx 280$  GeV along the diagonal. The latest CMS results for this simplified model are featured in [65–68,88] and exhibit a similar reach.

Further bottom squark decay modes have also been studied by ATLAS and CMS. For example, in a simplified model for the  $\tilde{b} \rightarrow t\tilde{\chi}_1^\pm$  decay mode, bottom squark quark masses below approximately 450 GeV are excluded [85,88,89].

The top squark decay modes depend on the SUSY mass spectrum, and on the  $\tilde{t}_L$ - $\tilde{t}_R$  mixture of the top squark mass eigenstate. If kinematically allowed, the two-body decays  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  (which requires  $m_{\tilde{t}} - m_{\tilde{\chi}_1^0} > m_t$ ) and  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$  (which requires  $m_{\tilde{t}} - m_{\tilde{\chi}_1^\pm} > m_b$ ) are expected to dominate. If not, the top squark decay may proceed either via the two-body decay  $\tilde{t} \rightarrow c\tilde{\chi}_1^0$  or through  $\tilde{t} \rightarrow bf\tilde{\chi}_1^0$  (where  $f$  and  $\bar{f}$  denote a fermion-antifermion pair with appropriate quantum numbers). For  $m_{\tilde{t}} - m_{\tilde{\chi}_1^0} > m_b$  the latter decay chain represents a four-body decay with a  $W$  boson, charged Higgs  $H$ , slepton  $\tilde{\ell}$ , or light flavor squark  $\tilde{q}$ , exchange. If the exchanged  $W$  boson and/or sleptons are kinematically allowed to be on-shell (



**Figure 6:** 95% C.L. exclusion contours in the sbottom-neutralino mass plane defined in the framework of a simplified model assuming a single decay chain of  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  [85].

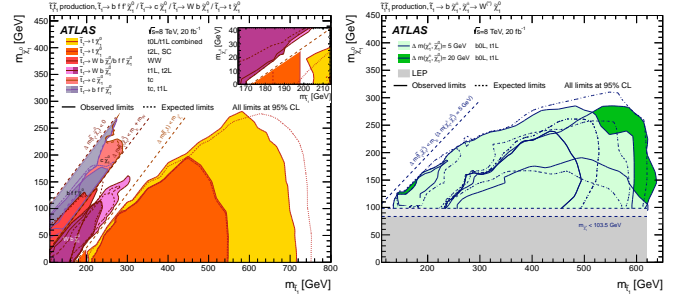
$m_{\tilde{t}} - m_{\tilde{\chi}_1^\pm} > m_b + m_W$  and/or  $m_{\tilde{t}} - m_{\tilde{t}} > m_b$ ), the three-body decays  $\tilde{t} \rightarrow Wb\tilde{\chi}_1^0$  and/or  $\tilde{t} \rightarrow b\tilde{\ell}\tilde{\nu}$  will become dominant. For further discussion on top squark decays see for example Ref. [90].

Limits from LEP on the  $\tilde{t}_1$  mass are  $m_{\tilde{t}} > 96$  GeV in the charm plus neutralino final state, and  $> 93$  GeV in the lepton, b-quark and sneutrino final state [61].

The Tevatron experiments have performed a number of searches for top squarks, often assuming direct pair production. In the  $b\tilde{\ell}\tilde{\nu}$  decay channel, and assuming a 100% branching fraction, limits are set as  $m_{\tilde{t}} > 210$  GeV for  $m_{\tilde{\nu}} < 110$  GeV and  $m_{\tilde{t}} - m_{\tilde{\nu}} > 30$  GeV, or  $m_{\tilde{t}} > 235$  GeV for  $m_{\tilde{\nu}} < 50$  GeV [91,92]. In the  $\tilde{t} \rightarrow c\tilde{\chi}_1^0$  decay mode, a top squark with a mass below 180 GeV is excluded for a neutralino lighter than 95 GeV [93,94]. In both analyses, no limits on the top squark can be set for heavy sneutrinos or neutralinos. In the  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$  decay channel, searches for a relatively light top squark have been performed in the dilepton final state [95,96]. The CDF experiment sets limits in the  $\tilde{t} - \tilde{\chi}_1^0$  mass plane for various branching fractions of the chargino decay to leptons and for two values of  $m_{\tilde{\chi}_1^\pm}$ . For  $m_{\tilde{\chi}_1^\pm} = 105.8$  GeV and  $m_{\tilde{\chi}_1^0} = 47.6$  GeV, top squarks between 128 and 135 GeV are excluded for  $W$ -like leptonic branching fractions of the chargino.

The LHC experiments have improved these limits substantially. As shown in the left plot of Fig. 7, limits on the top squark mass assuming a simplified model with a single decay chain of  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  reach up to about 700 GeV for light neutralinos, while for  $m_{\tilde{\chi}_1^0} > 270$  GeV no limits can be provided. The most important searches for this top squark decay topology are dedicated searches requiring zero or one isolated lepton, modest  $E_T^{\text{miss}}$ , and four or more jets out of which at least one jet must be reconstructed as  $b$ -jet [97,98]. To increase the sensitivity to this decay topology different signal regions are considered in these ATLAS analyses. Searches from the CMS collaboration

requiring one isolated lepton and using a boosted decision tree for a dedicated optimization in the  $m_{\tilde{t}} - m_{\tilde{\chi}_1^0}$  plane [99] or all-hadronic searches [65–68,88] provide comparable limits for this simplified model.



**Figure 7:** 95% C.L. exclusion contours in the  $m_{\tilde{t}} - m_{\tilde{\chi}_1^0}$  plane for different top squark decay chains and different searches from the ATLAS collaboration [85]. The left plot shows simplified model limits for three different decay chains;  $\tilde{t} \rightarrow c\tilde{\chi}_1^0$  ( $W$  and  $t$  forbidden),  $\tilde{t} \rightarrow Wb\tilde{\chi}_1^0$  ( $t$  forbidden), and  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  ( $t$  allowed), which represent three different kinematic regions of the top squark decay. The right plot shows simplified model limits for the top decay chain via a chargino:  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ ,  $\tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$  for different  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0}$ .

Assuming that the top squark decay exclusively proceeds via the chargino mediated decay chain  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ ,  $\tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$  yields stop mass exclusion limits that vary strongly with the assumptions made on the  $\tilde{t} - \tilde{\chi}_1^\pm - \tilde{\chi}_1^0$  mass hierarchy. As shown in the right plot of Fig. 7, above the universal chargino mass limit of  $m_{\tilde{\chi}_1^\pm} > 103.5$  GeV from LEP (see section II.5.1) the strongest limits are placed for nearly mass degenerate chargino and neutralinos. For  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} > 5$  GeV, a stop mass of  $\approx 650$  GeV for a light  $\tilde{\chi}_1^0$  is excluded, while no limit can be placed for  $m_{\tilde{\chi}_1^0} > 280$  GeV [85]. These limits, however, can weaken significantly when other assumptions about the mass hierarchy are imposed. For example, if the chargino becomes nearly mass degenerate with the top squark the key experimental signature turns from an all-hadronic final state with  $b$ -jets and  $E_T^{\text{miss}}$  into a multi-lepton and  $E_T^{\text{miss}}$  topology yielding significantly weaker limits for this top squark decay. As for the decay with top quarks in the final state, CMS [88,99] also provides comparable limits for this decay chain.

If the decays  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  and  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ ,  $\tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$  are kinematically forbidden, the decay chains  $\tilde{t} \rightarrow Wb\tilde{\chi}_1^0$  and  $\tilde{t} \rightarrow c\tilde{\chi}_1^0$  can become important. As shown in the left plot of Fig. 7, the one-lepton ATLAS search provides for the kinematic region  $m_{\tilde{t}} - m_{\tilde{\chi}_1^\pm} > m_b + m_W$  upper limits on top squark mass of  $\approx 300$  GeV for a neutralino lighter than  $\approx 170$  GeV [85], while the boosted decision tree based CMS analysis pushes this limit to about 320 GeV for neutralino masses below  $\approx 200$  GeV [99]. For the kinematic region in which even the production of

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real  $W$  bosons is not allowed, ATLAS and CMS improves the Tevatron limit on  $\tilde{t} \rightarrow c\tilde{\chi}^0$  substantially. Based on a combination of a monojet analysis and a dedicated charm quark identification algorithm, away from the kinematic boundary a top squark with a mass below 260 GeV is excluded by the ATLAS analysis for a neutralino lighter than 230 GeV [85]. Along the kinematic boundary for the  $\tilde{t} \rightarrow c\tilde{\chi}^0$  decay the ATLAS monojet results even excludes top squark masses below  $m_{\tilde{\chi}_1^0} \approx 260$  GeV. The corresponding CMS results [88] exhibit similar exclusions. The other decay chain relevant in this phase region is  $\tilde{t} \rightarrow b\tilde{f}\tilde{\chi}^0$ . Here the ATLAS one-lepton search [85] excludes up to  $m_{\tilde{t}} \approx 250$  GeV for  $m_{\tilde{\chi}_1^0}$  below 180 GeV, while the mono-jet excludes at the kinematic boundary top squarks below  $m_{\tilde{\chi}^\pm}$  below 260 GeV. Also for this decay chain CMS [88] provides similar results.

In general, the variety of top squark decay chains in the phase space region where  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  is kinematically forbidden represents a challenge for the experimental search program. As, for example, shown in the inlay of the left plot of Fig. 7 there are still regions in phase space for which the searches do not yet possess enough sensitivity to probe them. Additional data and more refined analyses will be required to also close these gaps.

R-parity violating production of single top squarks has been searched for at LEP, HERA, and the Tevatron. For example, an analysis from the ZEUS collaboration [100] makes an interpretation of its search result assuming top squarks to be produced via a  $\lambda'$  coupling and decay either to  $b\tilde{\chi}_1^\pm$  or R-parity-violating to a lepton and a jet. Limits are set on  $\lambda'_{131}$  as a function of the top squark mass in an MSSM framework with gaugino mass unification at the GUT scale. The search for top squark pair production in the context of R-parity violating supersymmetry has now also become a focus point for searches at the LHC. Recently the CMS collaboration has performed a search for top squarks using a variety of multilepton final states [101]. It provides lower limits on the top squark mass in models with non-zero leptonic R-parity violating couplings  $\lambda_{122}$  and  $\lambda_{233}$ . For a bino mass of 200 GeV, these limits are 1020 GeV and 820 GeV, respectively. The analysis also provides limits in a model with the semileptonic R-parity violating coupling  $\lambda'_{233}$ . The  $\lambda'$ -mediated top squark decay  $\tilde{t} \rightarrow b\ell$  has been studied by ATLAS for prompt decays [102], and by CMS for non-prompt decays [103]. The fully hadronic R-parity violating top squark decay  $\tilde{t} \rightarrow bs$ , involving  $\lambda''$ , has been searched for by ATLAS [104], and a lower top squark mass limit of 310 GeV was set in this decay mode. CMS [105] have searched for a top squark decay to a bottom quark and a light-flavor quark, and excludes top squarks with masses between 200 and 385 GeV in this decay mode.

Top squarks can also be long-lived and hadronize to a R-hadron, for example in the scenario where the top squark is the next-to-lightest SUSY particle (NLSP), with a small mass difference to the LSP. Searches for massive stable charged particles are sensitive to such top squarks. As shown in Fig. 4,

the CMS analysis [71] sets limits  $m_{\tilde{t}} > 800$  GeV in such scenarios, while ATLAS [72] reports limits of  $m_{\tilde{t}} > 900$  GeV. Limits from the Tevatron are about  $m_{\tilde{t}} > 300$  GeV [106,107].

It should be noted that limits discussed in this section belong to different top and bottom squark decay channels, different sparticle mass hierarchies, and different simplified decay scenarios. Therefore, care must be taken when interpreting these limits in the context of more complete SUSY models.

#### II.4.4. Summary of exclusion limits on squarks and gluinos assuming R-Parity conservation

A summary of the most important squark and gluino mass limits for different interpretation approaches assuming R-parity conservation is shown in Table 2.

For gluino masses rather similar limits, ranging from 1.2 TeV to 1.4 TeV, are obtained from different model assumptions indicating that the LHC is indeed probing for a large region in SUSY parameter space direct gluino production at the 1 TeV scale and beyond. However, for neutralino masses above approximately 700 GeV in the best case, ATLAS and CMS searches cannot place any limits on the gluino mass.

Limits on direct squark production, on the other hand, depend strongly on the chosen model. Especially for direct production of top squarks there are still large regions in parameter space where masses below 0.5 TeV cannot be excluded. This is also true for first and second generation squarks when only one single squark is considered. Furthermore, for neutralino masses above  $\approx 300$  GeV no limit on any direct squark production scenario can be placed by the LHC.

#### II.5. Exclusion limits on the masses of charginos and neutralinos

Charginos and neutralinos result from mixing of the charged wino and higgsino states, and the neutral bino, wino and higgsino states, respectively. The mixing is determined by a limited number of parameters. For charginos these are the wino mass parameter  $M_2$ , the higgsino mass parameter  $\mu$ , and  $\tan\beta$ , and for neutralinos these are the same parameters plus the bino mass parameter  $M_1$ . The mass states are four charginos  $\tilde{\chi}_1^\pm$ ,  $\tilde{\chi}_2^\pm$  and  $\tilde{\chi}_2^\pm$ , and four neutralinos  $\tilde{\chi}_1^0$ ,  $\tilde{\chi}_2^0$ ,  $\tilde{\chi}_3^0$  and  $\tilde{\chi}_4^0$ , ordered in increasing mass. The mass states are superpositions of the bino, wino and higgsino states. If any of the parameters  $M_1$ ,  $M_2$  or  $\mu$  happened to be substantially smaller than the others, the chargino and neutralino composition would be dominated by specific states, which are referred to as bino-like ( $M_1 \ll M_2, \mu$ ), wino-like ( $M_2 \ll M_1, \mu$ ), or higgsino-like ( $\mu \ll M_1, M_2$ ). If gaugino mass unification at the GUT scale is assumed, a relation between  $M_1$  and  $M_2$  at the electroweak scale follows:  $M_1 = 5/3 \tan^2 \theta_W M_2 \approx 0.5 M_2$ , with  $\theta_W$  the weak mixing angle. Charginos and neutralinos carry no color charge, and only have electroweak couplings (neglecting gravity).

##### II.5.1. Exclusion limits on chargino masses

If kinematically allowed, two body decay modes such as  $\tilde{\chi}^\pm \rightarrow \tilde{f}\tilde{f}'$  (including  $\ell\bar{\nu}$  and  $\bar{\ell}\nu$ ) are dominant. If not, three



**Table 2:** Summary of squark mass and gluino mass limits using different interpretation approaches assuming R-parity conservation. Masses in this table are provided in GeV. Further details about assumption and analyses from which these limits are obtained are discussed in the corresponding sections of the text.

Model	Assumption	$m_{\tilde{q}}$	$m_{\tilde{g}}$
CMSSM	$m_{\tilde{q}} \approx m_{\tilde{g}}$	$\approx 1700$	$\approx 1700$
	all $m_{\tilde{q}}$	-	$\approx 1300$
	all $m_{\tilde{g}}$	$\approx 1600$	-
Simplified model $\tilde{g}\tilde{q}, \tilde{g}\tilde{\bar{q}}$	$m_{\tilde{\chi}_1^0} = 0, m_{\tilde{q}} \approx m_{\tilde{g}}$	$\approx 1700$	$\approx 1700$
	$m_{\tilde{\chi}_1^0} = 0, \text{ all } m_{\tilde{q}}$	-	$\approx 1400$
	$m_{\tilde{\chi}_1^0} = 0, \text{ all } m_{\tilde{g}}$	$\approx 1400$	-
Simplified models $\tilde{g}\tilde{g}$			
$\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	-	$\approx 1300$
	$m_{\tilde{\chi}_1^0} > \approx 550$	-	no limit
$\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	-	$\approx 1350$
	$m_{\tilde{\chi}_1^0} > \approx 750$	-	no limit
$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	-	$\approx 1400$
	$m_{\tilde{\chi}_1^0} > \approx 700$	-	no limit
Simplified models $\tilde{q}\tilde{q}$			
$\tilde{q} \rightarrow q\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	$\approx 925$	-
	$m_{\tilde{\chi}_1^0} > \approx 350$	no limit	-
$\tilde{u}_L \rightarrow q\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	$\approx 575$	-
	$m_{\tilde{\chi}_1^0} > \approx 120$	no limit	-
$\tilde{b} \rightarrow b\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	$\approx 650$	-
	$m_{\tilde{\chi}_1^0} > \approx 280$	no limit	-
$\tilde{t} \rightarrow t\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} = 0$	$\approx 700$	-
	$m_{\tilde{\chi}_1^0} > \approx 270$	no limit	-
$\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$	$m_{\tilde{\chi}_1^0} = 0$	$\approx 650$	-
	$m_{\tilde{\chi}_1^0} > \approx 280$	no limit	-
$\tilde{t} \rightarrow Wb\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} < \approx 200$	$\approx 320$	-
$\tilde{t} \rightarrow c\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} < \approx 230$	$\approx 260$	-
	$m_{\tilde{t}} \approx m_{\tilde{\chi}_1^0}$	$\approx 260$	-
$\tilde{t} \rightarrow bf'f'\tilde{\chi}_1^0$	$m_{\tilde{\chi}_1^0} < \approx 180$	$\approx 250$	-
	$m_{\tilde{t}} \approx m_{\tilde{\chi}_1^0}$	$\approx 260$	-

body decay  $\tilde{\chi}^\pm \rightarrow f\bar{f}'\tilde{\chi}^0$  are mediated through virtual  $W$  bosons or sfermions. If sfermions are heavy, the  $W$  mediation dominates, and  $f\bar{f}'$  are distributed with branching fractions similar to  $W$  decay products (barring phase space effects for small mass gaps between  $\tilde{\chi}^\pm$  and  $\tilde{\chi}^0$ ). If, on the other hand, sleptons are light enough to play a significant role in the decay mediation, leptonic final states will be enhanced.

At LEP, charginos have been searched for in fully-hadronic, semi-leptonic and fully leptonic decay modes [108,109]. A general lower limit on the lightest chargino mass of 103.5 GeV

is derived, except in corners of phase space with low electron sneutrino mass, where destructive interference in chargino production, or two-body decay modes, play a role. The limit is also affected if the mass difference between  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_1^0$  is small; dedicated searches for such scenarios set a lower limit of 92 GeV.

At the Tevatron, charginos have been searched for via associated production of  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  [110,111]. Decay modes involving multilepton final states provide the best discrimination against the large multijet background. Analyses have looked for at least three charged isolated leptons, for two leptons with missing transverse momentum, or for two leptons with the same charge. Depending on the  $(\tilde{\chi}_1^\pm - \tilde{\chi}_1^0)$  and/or  $(\tilde{\chi}_2^0 - \tilde{\chi}_1^0)$  mass differences, leptons may be soft.

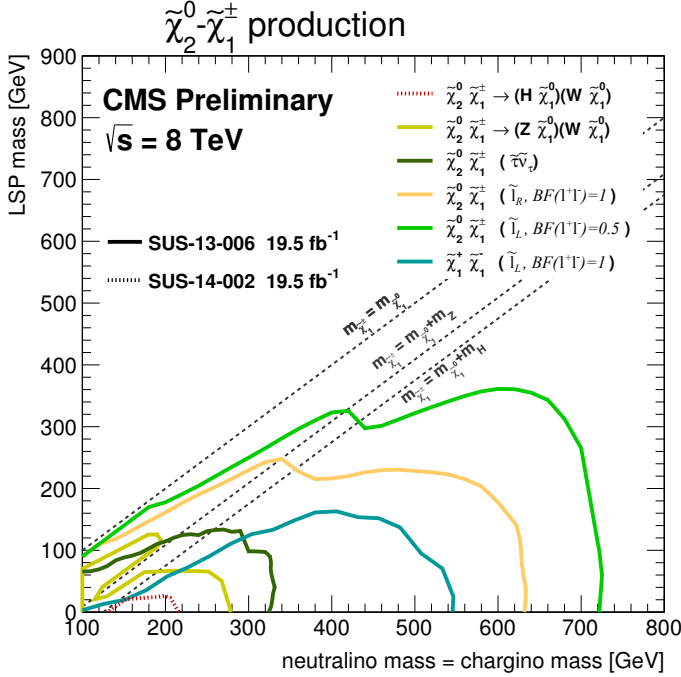
At the LHC, the search strategy is similar to that at the Tevatron. As shown in Fig. 1, the cross section of pair production of chargino and neutralinos at the LHC, for masses of several hundreds of GeV, is at least two orders of magnitude smaller than for colored SUSY particles (e.g. top squark pair production). For this reason a high statistics data sample is required to improve the sensitivity of LEP and Tevatron searches for direct chargino/neutralino production. With the full LHC Run 1 data, ATLAS and CMS have started to surpass the limits from LEP and Tevatron in regions of SUSY parameter space.

Chargino pair production is searched for in the dilepton plus missing momentum final state. In the simplified model interpretation of the results, assuming mediation of the chargino decay by light sleptons, ATLAS [112] and CMS [113] set limits on the chargino mass up to 540 GeV for massless LSPs, but no limits on the chargino mass can be set for  $\tilde{\chi}_1^0$  heavier than 180 GeV. Limits are fairly robust against variation of the slepton mass, unless the mass gap between chargino and slepton becomes small. First limits are also set on charginos decaying via a  $W$  boson [114]: chargino masses below 180 GeV are excluded for massless LSPs, but no limits are set for LSPs heavier than 25 GeV. The trilepton plus missing momentum final state is used to set limits on  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  production, assuming wino-like  $\tilde{\chi}^\pm$  and  $\tilde{\chi}_2^0$ , bino-like  $\tilde{\chi}_1^0$ , and  $m_{\tilde{\chi}^\pm} = m_{\tilde{\chi}_2^0}$ , leaving  $m_{\tilde{\chi}^\pm}$  and  $m_{\tilde{\chi}_1^0}$  free. Again, the branching fraction of leptonic final states is determined by the slepton masses. If the decay is predominantly mediated by a light  $\tilde{\ell}_L$ , i.e.  $\tilde{\ell}_R$  is assumed to be heavy, the three lepton flavors will be produced in equal amounts. It is assumed that  $\tilde{\ell}_L$  and sneutrino masses are equal, and diagrams with sneutrinos are included. In this scenario, ATLAS [112] and CMS [113] exclude chargino masses below 730 GeV for massless LSPs; no limits are set for LSP masses above 400 GeV. If the decay is dominated by a light  $\tilde{\ell}_R$ , the chargino cannot be a pure wino but needs to have a large higgsino component, preferring the decays to tau leptons. Limits are set in various scenarios. If, like for  $\tilde{\ell}_L$ , a flavor-democratic scenario is assumed, CMS sets limits of 620 GeV on the chargino mass for massless LSPs, but under the assumption that both  $\tilde{\chi}^\pm$  and  $\tilde{\chi}_2^0$  decay leads to tau leptons in the final

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state, the chargino mass limit deteriorates to 350 GeV for massless LSPs [113]. ATLAS assumes a simplified model in which staus are significantly lighter than the other sleptons in order to search for a similar multi-tau final state, and sets a similar limit on the chargino mass [112].



**Figure 8:** A summary of limits on chargino and neutralino masses in simplified models as obtained by CMS [158].

If sleptons are heavy, the chargino is assumed to decay to a  $W$  boson plus LSP, and the  $\tilde{\chi}_2^0$  into  $Z$  plus LSP or  $H$  plus LSP. In the  $WZ$  channel, ATLAS [112] and CMS [113] limits on the chargino mass reach 420 GeV for massless LSPs, but no limits are set for LSPs heavier than 150 GeV. In the  $WH$  channel, for  $m_H = 125$  GeV and using Higgs decays to  $b\bar{b}$ ,  $\gamma\gamma$  and  $WW$  (ATLAS [115]), or Higgs decays to  $b\bar{b}$ ,  $\gamma\gamma$ ,  $WW$ ,  $ZZ$  and  $\tau^+\tau^-$  (CMS [113,116]), assuming a SM-like branching fraction in these final states, chargino mass limits extend up to 250 GeV for massless LSPs, but vanish for LSPs above 40 GeV.

The CMS results on electroweak gaugino searches interpreted in simplified models are summarized in Fig. 8, the ATLAS results are similar.

In both the wino region (a characteristic of anomaly-mediated SUSY breaking models) and the higgsino region of the MSSM, the mass splitting between  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_1^0$  is small. The chargino decay products are very soft and may escape detection. These compressed spectra are very hard to find, and have triggered dedicated search strategies, which, however, still have limited sensitivity. Photons or jets from initial state radiation may be used to tag such decays. An alternative production

mode of electroweak gauginos is provided by vector-boson-fusion, where two additional jets with a large rapidity gap can be used to select events and suppress backgrounds [112,117].

In scenarios with compressed spectra, charginos may be long-lived. Charginos decaying in the detectors away from the primary vertex could lead to signatures such as kinked-tracks, or apparently disappearing tracks, since, for example, the pion in  $\tilde{\chi}_1^\pm \rightarrow \pi^\pm \tilde{\chi}_1^0$  might be too soft to be reconstructed. At the LHC, searches have been performed for such disappearing tracks, and interpreted within anomaly-mediated SUSY breaking models [118,119]. Charginos with lifetimes between 0.1 and 10 ns are excluded for chargino masses up to 500 GeV. Within AMSB models, a lower limit on the chargino mass of 270 GeV is set, for a mass difference with the LSP of 160 MeV and a lifetime of 0.2 ns.

Charginos with a lifetime longer than the time needed to pass through the detector appear as charged stable massive particles. Limits have been derived by the LEP experiments [120], by D0 at the Tevatron [107], and by the LHC experiments [72,121,122]. For lifetimes above 100 ns, charginos below some 800 GeV are excluded.

#### II.5.2. Exclusion limits on neutralino masses

In a considerable part of the MSSM parameter space, and in particular when demanding that the LSP carries no electric or color charge, the lightest neutralino  $\tilde{\chi}_1^0$  is the LSP. If R-parity is conserved, such a  $\tilde{\chi}_1^0$  is stable. Since it is weakly interacting, it will typically escape detectors unseen. Limits on the invisible width of the  $Z$  boson apply to neutralinos with a mass below 45.5 GeV, but depend on the  $Z$ -neutralino coupling. Such a coupling could be small or even absent; in such a scenario there is no general lower limit on the mass of the lightest neutralino [123]. In models with gaugino mass unification and sfermion mass unification at the GUT scale, a lower limit on the neutralino mass is derived from limits from direct searches, notably for charginos and sleptons, and amounts to 47 GeV [124]. Assuming a constraining model like the CMSSM, this limit increases to 50 GeV at LEP; however the strong constraints now set by the LHC increase such CMSSM-derived  $\tilde{\chi}_1^0$  mass limits to well above 200 GeV [125].

In gauge-mediated models, the LSP is typically a gravitino, and the phenomenology is determined by the nature of the next-to-lightest supersymmetric particle (NLSP). A NLSP neutralino will decay to a gravitino and a SM particle whose nature is determined by the neutralino composition. Final states with two high  $p_T$  photons and missing momentum are searched for, and interpreted in gauge mediation models with bino-like neutralinos [126–130].

Assuming the production of at least two neutralinos per event, neutralinos with large non-bino components can also be searched for by their decay in final states with missing momentum plus any two bosons out of the collection  $\gamma, Z, H$ . A number of searches at the LHC have tried to cover the rich phenomenology of the various  $Z$  and  $H$  decay modes [113,116,129–131].

In gauge mediation models, NLSP neutralino decays need not be prompt, and experiments have searched for late decays with photons in the final state. CDF have searched for delayed  $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$  decays using the timing of photon signals in the calorimeter [132]. CMS has used the same technique at the LHC [133]. Results are given as upper limits on the neutralino production cross section as a function of neutralino mass and lifetime. D0 has looked at the direction of showers in the electromagnetic calorimeter with a similar goal [134], and ATLAS has searched for photon candidates that do not point back to the primary vertex, as well as for delayed photons [135].

Heavier neutralinos, in particular  $\tilde{\chi}_2^0$ , have been searched for in their decays to the lightest neutralino plus a  $\gamma$ , a  $Z$  boson or a Higgs boson. Limits on electroweak production of  $\tilde{\chi}_2^0$  plus  $\tilde{\chi}_1^\pm$  from trilepton analyses have been discussed in the section on charginos; the assumption of equal mass of  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_1^\pm$  make the limits on chargino masses apply to  $\tilde{\chi}_2^0$  as well. Multilepton analyses have also been used to set limits on  $\tilde{\chi}_2^0 \tilde{\chi}_3^0$  production; assuming equal mass limits are set up to 680 GeV for massless LSPs [112]. Again, compressed spectra with small mass differences between the heavier neutralinos and the LSP form the most challenging region.

In  $\tilde{\chi}_2^0$  decays to  $\tilde{\chi}_1^0$  and a lepton pair, the lepton pair invariant mass distribution may show a structure that can be used to measure the  $\tilde{\chi}_2^0 - \tilde{\chi}_1^0$  mass difference in case of a signal [34]. This structure, however, can also be used in the search strategy itself, as demonstrated by CMS [136] and ATLAS [131].

In model with R-parity violation, the lightest neutralino can decay even if it is the lightest supersymmetric particle. If the decay involves a non-zero  $\lambda$  coupling, the final state will be a multi-lepton one. Searches for events with four or more isolated charged leptons by ATLAS [77,137] and CMS [78] are interpreted in such models. With very small coupling values, the neutralino would be long-lived, leading to lepton pairs with a displaced vertex, which have also been searched for [73,84]. Searches for events with a displaced hadronic vertex, with or without a matched lepton, are interpreted in a model with R-parity violating neutralino decay involving a non-zero  $\lambda'$  coupling [73,138]. Neutralino decays involving non-zero  $\lambda''$  lead to fully hadronic final states, and searches for jet-pair resonances are used to set limits, typically on the production of colored particles like top squarks or gluinos, which are assumed to be the primary produced sparticles in these interpretations, as discussed earlier.

Interpretations of the search results outside simplified models, such as in the phenomenological MSSM [139–141], show that the simplified model limits must be interpreted with care. Electroweak gauginos in models that are compatible with the relic density of dark matter in the universe, for example, have particularly tuned mixing parameters and mass spectra, which are not always captured by the simplified models used.

**Table 3:** Summary of weak gaugino mass limits in simplified models, assuming R-parity conservation. Masses in the table are provided in GeV. Further details about assumptions and analyses from which these limits are obtained are discussed in the text.

Assumption	$m_\chi$
$\tilde{\chi}_1^\pm$ , all $\Delta m(\tilde{\chi}_1^\pm, \tilde{\chi}_1^0)$	$> 92$
$\tilde{\chi}_1^\pm$ $\Delta m > 5$ , $m_{\tilde{\nu}} > 300$	$> 103.5$
$\tilde{\chi}_1^\pm$ , $m_{(\tilde{\ell}, \tilde{\nu})} = (m_{\tilde{\chi}_1^\pm} + m_{\tilde{\chi}_1^0})/2$ $m_{\tilde{\chi}_1^0} \approx 0$	$> 540$
$\tilde{\chi}_1^\pm$ , $m_{\tilde{\chi}_1^0} > 180$	no LHC limit
$\tilde{\chi}_1^\pm$ , $m_{\tilde{\ell}} > m_{\tilde{\chi}_1^\pm}$ $m_{\tilde{\chi}_1^0} \approx 0$	$> 180$
$\tilde{\chi}_1^\pm$ , $m_{\tilde{\chi}_1^0} > 25$	no LHC limit
$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\ell}_L} = (m_{\tilde{\chi}_1^\pm} + m_{\tilde{\chi}_1^0})/2$ $m_{\tilde{\chi}_1^0} \approx 0$	$> 730$
$m_{\tilde{\chi}_1^0} > 400$	no LHC limit
$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\ell}_R} = (m_{\tilde{\chi}_1^\pm} + m_{\tilde{\chi}_1^0})/2$ $m_{\tilde{\chi}_1^0} \approx 0$	flavor-democratic $> 620$
$m_{\tilde{\chi}_1^0} > 220$	no LHC limit
$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\tau}} = (m_{\tilde{\chi}_1^\pm} + m_{\tilde{\chi}_1^0})/2$ $m_{\tilde{\chi}_1^0} \approx 0$	$\tilde{\tau}$ -dominated $> 350$
$m_{\tilde{\chi}_1^0} > 120$	no LHC limit
$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\ell}} > m_{\tilde{\chi}_1^\pm}$ , $\text{BF}(WZ) = 1$ $m_{\tilde{\chi}_1^0} \approx 0$	$> 420$
$m_{\tilde{\chi}_1^0} > 150$	no LHC limit
$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\ell}} > m_{\tilde{\chi}_1^\pm}$ , $\text{BF}(WH) = 1$ $m_{\tilde{\chi}_1^0} \approx 0$	$> 250$
$m_{\tilde{\chi}_1^0} > 40$	no LHC limit

## II.6. Exclusion limits on slepton masses

In models with slepton and gaugino mass unification at the GUT scale, the right-handed slepton,  $\tilde{\ell}_R$ , is expected to be lighter than the left-handed slepton,  $\tilde{\ell}_L$ . For tau sleptons there may be considerable mixing between the L and R states, leading to a significant mass difference between the lighter  $\tilde{\tau}_1$  and the heavier  $\tilde{\tau}_2$ .

### II.6.1. Exclusion limits on the masses of charged sleptons

The most model-independent searches for selectrons, smuons and staus originate from the LEP experiments [142]. Smuon production only takes place via s-channel  $\gamma^*/Z$  exchange. Search results are often quoted for  $\tilde{\mu}_R$ , since it is typically lighter than  $\tilde{\mu}_L$  and has a weaker coupling to the  $Z$  boson; limits are therefore conservative. Decays are expected to be dominated by  $\tilde{\mu}_R \rightarrow \mu \tilde{\chi}_1^0$ , leading to two non-back-to-back muons and missing momentum. Slepton mass limits are calculated in the MSSM under the assumption of gaugino mass unification at the GUT scale, and depend on the mass difference

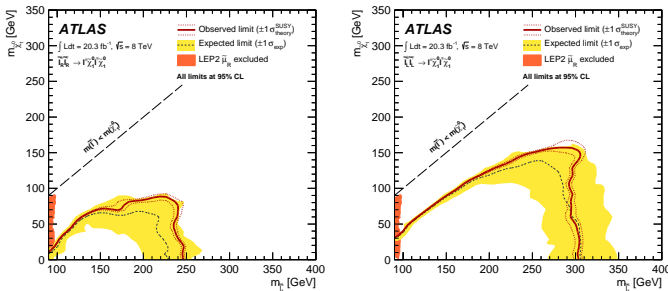
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between the smuon and  $\tilde{\chi}_1^0$ . A  $\tilde{\mu}_R$  with a mass below 94 GeV is excluded for  $m_{\tilde{\mu}_R} - m_{\tilde{\chi}_1^0} > 10$  GeV. The selectron case is similar to the smuon case, except that an additional production mechanism is provided by t-channel neutralino exchange. The  $\tilde{e}_R$  lower mass limit is 100 GeV for  $m_{\tilde{\chi}_1^0} < 85$  GeV. Due to the t-channel neutralino exchange,  $\tilde{e}_R \tilde{e}_L$  pair production was possible at LEP, and a lower limit of 73 GeV was set on the selectron mass regardless of the neutralino mass by scanning over MSSM parameter space [143]. The potentially large mixing between  $\tilde{\tau}_L$  and  $\tilde{\tau}_R$  not only makes the  $\tilde{\tau}_1$  light, but can also make its coupling to the Z boson small. LEP lower limits on the  $\tilde{\tau}$  mass range between 87 and 93 GeV depending on the  $\tilde{\chi}_1^0$  mass, for  $m_{\tilde{\tau}} - m_{\tilde{\chi}_1^0} > 7$  GeV [142].

As shown in Fig. 1, at the LHC pair production of sleptons is not only heavily suppressed with respect to pair production of colored SUSY particles but the cross section is also almost two orders of magnitude smaller than the one of pair production of chargino and neutralinos. Only with the full Run 1 LHC data set, ATLAS and CMS have started to surpass the sensitivity of the LEP analyses under certain assumptions.

ATLAS and CMS have searched for direct production of selectron pairs and smuon pairs at the LHC, with each slepton decaying to its corresponding SM partner lepton and the  $\tilde{\chi}_1^0$  LSP. ATLAS [114] and CMS [113] set limits in this model of 240 GeV for  $\tilde{\ell}_R$ , and 290 GeV for  $\tilde{\ell}_L$ , for a massless  $\tilde{\chi}_1^0$  and assuming equal selectron and smuon masses, as shown in Fig. 9. The limits deteriorate with increasing  $\tilde{\chi}_1^0$  mass due to decreasing missing momentum and lepton momentum. As a consequence, there is a gap between LEP and LHC limits for  $\tilde{\chi}_1^0$  masses above 20 GeV, and no limits are set for  $\tilde{\chi}_1^0$  masses above 90 GeV ( $\tilde{\ell}_R$ ) or above 150 GeV ( $\tilde{\ell}_L$ ).



**Figure 9:** Exclusion limits on  $\tilde{\ell}_R$  masses (left) and  $\tilde{\ell}_L$  masses (right), assuming equal selectron and smuon masses in both scenarios, and assuming a 100% branching fraction for  $\tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$  [114].

In gauge-mediated SUSY breaking models, sleptons can be (co-)NLSPs, *i.e.*, the next-to-lightest SUSY particles and almost degenerate in mass, decaying to a lepton and a gravitino. This decay can either be prompt, or the slepton can have a non-zero lifetime. Combining several analyses, lower mass limits on  $\tilde{\mu}_R$  of 96.3 GeV and on  $\tilde{e}_R$  of 66 GeV are set for all

slepton lifetimes at LEP [144]. In a considerable part of parameter space in these models, the  $\tilde{\tau}$  is the NLSP. The LEP experiments have set lower limits on the mass of such a  $\tilde{\tau}$  between 87 and 97 GeV, depending on the  $\tilde{\tau}$  lifetime. ATLAS has searched for final states with  $\tau$ s, jets and missing transverse momentum, and has interpreted the results in GMSB models setting limits on the model parameters [145]. CMS has interpreted a multilepton analysis in terms of limits on gauge mediation models with slepton (co-)NLSP [146]. CDF has put limits on gauge mediation models at high  $\tan\beta$  and slepton (co-)NLSP using an analysis searching for like-charge light leptons and taus [147].

Limits also exist on sleptons in R-parity violating models, both from LEP and the Tevatron experiments. From LEP, lower limits on  $\tilde{\mu}_R$  and  $\tilde{e}_R$  masses in such models are 97 GeV, and the limits on the stau mass are very close: 96 GeV [148].

Charged slepton decays may be kinematically suppressed, for example in the scenario of a NLSP slepton with a very small mass difference to the LSP. Such a slepton may appear to be a stable charged massive particle. Interpretation of searches at LEP for such signatures within GMSB models with stau NLSP or slepton co-NLSP exclude masses up to 99 GeV [120]. Searches of stable charged particles at the Tevatron [106,107] and at the LHC [71,72] are also interpreted in terms of limits on stable charged sleptons. The limits obtained at the LHC exclude stable staus with masses below 339 GeV when produced directly in pairs, and below 500 GeV when staus are produced both directly and indirectly in the decay of other particles in a GMSB model.

**Table 4:** Summary of slepton mass limits from LEP and LHC, assuming R-parity conservation and 100% branching fraction for  $\tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$ . Masses in this table are provided in GeV.

Assumption	$m_{\tilde{\ell}}$
$\tilde{\mu}_R, \Delta m(\tilde{\mu}_R, \tilde{\chi}_1^0) > 10$	$> 94$
$\tilde{e}_R, \Delta m(\tilde{e}_R, \tilde{\chi}_1^0) > 10$	$> 94$
$\tilde{e}_R$ , any $\Delta m$	$> 73$
$\tilde{\tau}_R, \Delta m((\tilde{\tau}_R, \tilde{\chi}_1^0) > 7$	$> 87$
$\tilde{\nu}_e, \Delta m(\tilde{e}_R, \tilde{\chi}_1^0) > 10$	$> 94$
$m_{\tilde{e}_R} = m_{\tilde{\mu}_R}, m_{\tilde{\chi}_1^0} \approx 0$	$> 240$
$m_{\tilde{\chi}_1^0} > \approx 90$	no LHC limit
$m_{\tilde{e}_L} = m_{\tilde{\mu}_L}, m_{\tilde{\chi}_1^0} \approx 0$	$> 290$
$m_{\tilde{\chi}_1^0} > \approx 150$	no LHC limit

### II.6.2. Exclusion limits on sneutrino masses

The invisible width of the Z boson puts a lower limit on the sneutrino mass of about 45 GeV. Tighter limits are derived from other searches, notably for gauginos and sleptons, under the assumption of gaugino and sfermion mass universality at the GUT scale, and amount to approximately 94 GeV in the MSSM [149]. It is possible that the lightest sneutrino is the

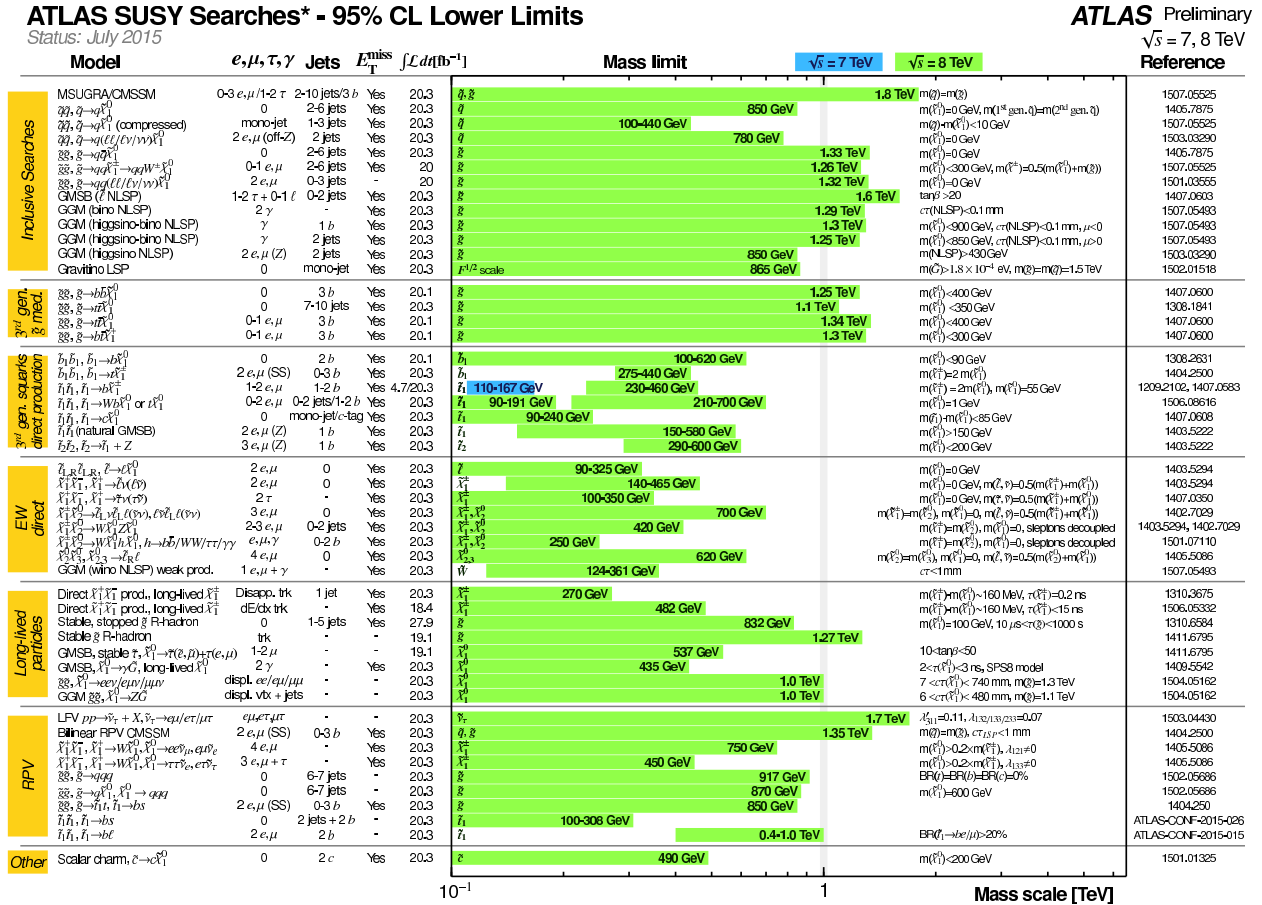
LSP; however, a left-handed sneutrino LSP is ruled out as a cold dark matter candidate [150,151].

Production of pairs of sneutrinos in R-parity violating models has been searched for at LEP [148]. Assuming fully leptonic decays via  $\lambda$ -type couplings, lower mass limits between 85 and 100 GeV are set. At the Tevatron [152,153] and at the LHC [154,155], searches have focused on scenarios with resonant production of a sneutrino, decaying to  $e\mu$ ,  $\mu\tau$  and  $e\tau$  final states. No signal has been seen, and limits have been set on sneutrino masses as a function of the value of relevant RPV couplings. As an example, the LHC experiments exclude a resonant tau sneutrino with a mass below 1500 GeV for  $\lambda_{312} > 0.07$  and  $\lambda'_{311} > 0.01$ .

experiments, flavor physics, high-precision electroweak results, and astrophysical data.

In the pre-LHC era these fits were mainly dominated by indirect constraints. Even for very constrained models like the CMSSM, the allowed parameter space, in terms of squark and gluino masses, ranged from several hundreds of GeV to a few TeV. Furthermore, these global fits indicated that squarks and gluino masses in the range of 500 to 1000 GeV were the preferred region of parameter space, although values as high as few TeV were allowed with lower probabilities [156].

With ATLAS and CMS now probing mass scales around 1 TeV and even beyond, the importance of the direct searches for global analyses of allowed SUSY parameter space has strongly



\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus  $1\sigma$  theoretical signal cross section uncertainty.

**Figure 10.** Overview of the current landscape of SUSY searches at the LHC. The plot shows exclusion mass limits of ATLAS for different searches and interpretation assumptions [157]. The corresponding results of CMS are comparable.

## II.7. Global interpretations

Apart from the interpretation of direct searches for sparticle production at colliders in terms of limits on masses of individual SUSY particles, model-dependent interpretations of allowed SUSY parameter space are derived from global SUSY fits. Typically these fits combine the results from collider experiments with indirect constraints on SUSY as obtained from low-energy

increased. For example, imposing the new experimental limits on constrained supergravity models pushes the most likely values of first generation squark and gluino masses significantly beyond 1 TeV, typically resulting in overall values of fit quality much worse than those in the pre-LHC era [125]. Although these constrained models are not yet ruled out, the extended

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experimental limits impose tight constraints on the allowed parameter space.

For this reason, the emphasis of global SUSY fits has shifted towards less-constrained SUSY models. Especially interpretations in the pMSSM [121,139–141] but also in simplified models have been useful to generalize SUSY searches, for example to redesign experimental analyses in order to increase their sensitivity for compressed spectra, where the mass of the LSP is much closer to squark and gluino masses than predicted, for example, by the CMSSM. As shown in Table 2, for neutralino masses above a few hundred GeV the current set of ATLAS and CMS searches cannot exclude the existence of light squarks and also gluinos above approximately 1 TeV are not yet fully excluded.

Furthermore, the discovery of a Higgs boson with a mass around 125 GeV has triggered many studies regarding the compatibility of SUSY parameter space with this new particle. Much of it is still work in progress and it will be interesting to see how the interplay between the results from direct SUSY searches and more precise measurements of the properties of the Higgs boson will unfold in the forthcoming era of high-energy running of the LHC.

### II.8. Summary and Outlook

Direct searches for SUSY, combined with limits from high-precision experiments that look for new physics in loops, put SUSY under considerable scrutiny. In particular the absence of any observation of new phenomena at the first run of the LHC, at  $\sqrt{s} = 7$  and 8 TeV, place significant constraints on SUSY parameter space. Today, inclusive searches probe production of gluinos in the range of 1.0 – 1.4 TeV, first and second generation squarks to about 1.0 TeV, third generation squarks at scales around 600 GeV, electroweak gauginos at scales around 300 – 500 GeV, and sleptons around 200 GeV. However, depending on the assumptions made of the underlying SUSY spectrum these limits can also weaken considerably. An overview of the current landscape of SUSY searches and corresponding exclusion limits at the LHC is shown in Fig. 10 from the ATLAS experiment [157]. The corresponding results of the CMS experiment are similar [158].

The interpretation of results at the LHC has moved away from constrained models like the CMSSM towards a large set of simplified models, or the pMSSM. On the one hand this move is because the LHC limits have put constrained models like the CMSSM under severe pressure, while on the other hand simplified models leave more freedom to vary parameters and form a better representation of the underlying sensitivity of analyses. However, these interpretations in simplified models do not come without a price: the decomposition of a potentially complicated reality in a limited set of individual decay chains can be significantly incomplete. Therefore, quoted limits in simplified models are only valid under the explicit assumptions made in these models, assumptions that are usually stated on the plots, and in the relevant LHC papers. Interpretations of simplified models in generic cases, ignoring the assumptions made, can

lead to overestimation of limits on SUSY parameter space. The recent addition of more comprehensive interpretations in the pMSSM is expected to overcome some of the limitations arising from the characterisation of searches in simplified model and thus will enable an even more refined understanding of the probed SUSY parameter space. In this context, the limit range of 1.0 – 1.4 TeV on generic colored SUSY particles only holds for light neutralinos, in the R-parity conserving MSSM. Limits on third generation squarks and electroweak gauginos also only hold for light neutralinos, and under specific assumptions for decay modes and slepton masses. In general, SUSY below the 1 TeV scale is not yet ruled out.

The new LHC run at  $\sqrt{s} = 13$  TeV, with significantly larger integrated luminosities, will present again a great opportunity for SUSY searches. The operation at higher energy will increase the production cross section for SUSY particles, shown in Fig. 1, substantially. While typically for masses around 500 GeV the increase is about 3 to 5 times the production cross section at 8 TeV, this becomes an increase of almost two orders of magnitude for a SUSY mass scale of 1.5 to 2 TeV. Apart from pushing the sensitivity of LHC searches to higher mass scales, further LHC data will also help to reduce holes and gaps that are left behind in today's SUSY limits. These could be, for example, due to compressed particle spectra, stealth SUSY, or the violation of R-parity.

### References

1. H. Miyazawa, *Prog. Theor. Phys.* **36**, 1266 (1966).
2. Yu. A. Golfand and E.P. Likhtman, *Sov. Phys. JETP Lett.* **13**, 323 (1971).
3. J.L. Gervais and B. Sakita, *Nucl. Phys.* **B34**, 632 (1971).
4. D.V. Volkov and V.P. Akulov, *Phys. Lett.* **B46**, 109 (1973).
5. J. Wess and B. Zumino, *Phys. Lett.* **B49**, 52 (1974).
6. J. Wess and B. Zumino, *Nucl. Phys.* **B70**, 39 (1974).
7. A. Salam and J.A. Strathdee, *Nucl. Phys.* **B76**, 477 (1974).
8. H.P. Nilles, *Phys. Reports* **110**, 1 (1984).
9. H.E. Haber and G.L. Kane, *Phys. Reports* **117**, 75 (1987).
10. E. Witten, *Nucl. Phys.* **B188**, 513 (1981).
11. S. Dimopoulos and H. Georgi, *Nucl. Phys.* **B193**, 150 (1981).
12. M. Dine, W. Fischler, and M. Srednicki, *Nucl. Phys.* **B189**, 575 (1981).
13. S. Dimopoulos and S. Raby, *Nucl. Phys.* **B192**, 353 (1981).
14. N. Sakai, *Z. Phys.* **C11**, 153 (1981).
15. R.K. Kaul and P. Majumdar, *Nucl. Phys.* **B199**, 36 (1982).
16. H. Goldberg, *Phys. Rev. Lett.* **50**, 1419 (1983).
17. J.R. Ellis *et al.*, *Nucl. Phys.* **B238**, 453 (1984).
18. G. Jungman and M. Kamionkowski, *Phys. Reports* **267**, 195 (1996).
19. S. Dimopoulos, S. Raby, and F. Wilczek, *Phys. Rev. D* **24**, 1681 (1981).

20. W.J. Marciano and G. Senjanović, Phys. Rev. **D25**, 3092 (1982).
21. M.B. Einhorn and D.R.T. Jones, Nucl. Phys. **B196**, 475 (1982).
22. L.E. Ibanez and G.G. Ross, Phys. Lett. **B105**, 439 (1981).
23. N. Sakai, Z. Phys. **C11**, 153 (1981).
24. U. Amaldi, W. de Boer, and H. Furstenau, Phys. Lett. **B260**, 447 (1991).
25. P. Langacker and N. Polonsky, Phys. Rev. **D52**, 3081 (1995).
26. P. Fayet, Phys. Lett. **B64**, 159 (1976).
27. G.R. Farrar and P. Fayet, Phys. Lett. **B76**, 575 (1978).
28. H.E. Haber, *Supersymmetry, Part I (Theory)*, in this Review.
29. CMS Collab. and LHCb Collab., Nature **522**, 68 (2015).
30. G. Hinshaw *et al.*, Astrophys. J. Supp. **208**, 19H (2013).
31. Planck Collab, *Planck 2015 results. XIII. Cosmological parameters*, arXiv:1502.01589 [astro-ph.CO].
32. M. Carena *et al.*, *Higgs Bosons: Theory and Searches*, this Review.
33. J.F. Grivaz, *Supersymmetry, Part II (Experiment)*, in: 2010 Review of Particle Physics, K. Nakamura *et al.*, (Particle Data Group), J. Phys. **G37**, 075021 (2010).
34. I. Hinchliffe *et al.*, Phys. Rev. **D55**, 5520 (1997).
35. L. Randall and D. Tucker-Smith, Phys. Rev. Lett. **101**, 221803 (2008).
36. CMS Collab., Phys. Lett. **B698**, 196 (2011).
37. CMS Collab., Phys. Rev. Lett. **107**, 221804 (2011).
38. CMS Collab., JHEP **1301**, 077 (2013).
39. CMS Collab., Eur. Phys. J. **C73**, 2568 (2013).
40. CMS Collab., Phys. Rev. **D85**, 012004 (2012).
41. C.G. Lester and D.J. Summers, Phys. Lett. **B463**, 99 (1999).
42. D.R. Tovey, JHEP **0804**, 034 (2008).
43. A.H. Chamseddine, R. Arnowitt, and P. Nath, Phys. Rev. Lett. **49**, 970 (1982).
44. E. Cremmer *et al.*, Nucl. Phys. **B212**, 413 (1983).
45. P. Fayet, Phys. Lett. **B70**, 461 (1977).
46. M. Dine, A.E. Nelson, and Yu. Shirman, Phys. Rev. **D51**, 1362 (1995).
47. G.F. Giudice *et al.*, JHEP **9812**, 027 (1998).
48. L. Randall and R. Sundrum, Nucl. Phys. **B557**, 79 (1999).
49. R. Arnowitt and P. Nath, Phys. Rev. Lett. **69**, 725 (1992).
50. G.L. Kane *et al.*, Phys. Rev. **D49**, 6173 (1994).
51. P. Meade, N. Seiberg, and D. Shih, Prog. Theor. Phys. Supp. **177**, 143 (2009).
52. W. Beenakker *et al.*, Nucl. Phys. **B492**, 51 (1997); W. Beenakker *et al.*, Nucl. Phys. **B515**, 3 (1998); W. Beenakker *et al.*, Phys. Rev. Lett. **83**, 3780 (1999), Erratum *ibid.*, **100**, 029901 (2008); M. Spira, hep-ph/0211145 (2002); T. Plehn, Czech. J. Phys. **55**, B213 (2005).
53. A. Djouadi, J-L. Kneur, and G. Moultaka, Comp. Phys. Comm. **176**, 426 (2007).
54. C.F. Berger *et al.*, JHEP **0902**, 023 (2009).
55. H. Baer *et al.*, hep-ph/9305342, 1993.
56. R.M. Barnett, H.E. Haber, and G.L. Kane, Nucl. Phys. **B267**, 625 (1986).
57. H. Baer, D. Karatas, and X. Tata, Phys. Lett. **B183**, 220 (1987).
58. J. Alwall, Ph.C. Schuster, and N. Toro, Phys. Rev. **D79**, 075020 (2009).
59. J. Alwall *et al.*, Phys. Rev. **D79**, 015005 (2009).
60. O. Buchmueller and J. Marrouche, Int. J. Mod. Phys. **A29**, 1450032 (2014).
61. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/04-02.1, <http://lepsusy.web.cern.ch/lepsusy>.
62. ATLAS Collab., JHEP **1510**, 054 (2015).
63. CDF Collab., Phys. Rev. Lett. **102**, 121801 (2009).
64. D0 Collab., Phys. Lett. **B660**, 449 (2008).
65. CMS Collab., JHEP **1505**, 078 (2015).
66. CMS Collab., Phys. Lett. **B725**, 243 (2013).
67. CMS Collab., Eur. Phys. J. **C73**, 2568 (2013).
68. CMS Collab., Phys. Rev. **D91**, 052018 (2015).
69. ATLAS Collab., JHEP **1410**, 24 (2014).
70. CMS Collab., Phys. Lett. **B733**, 328 (2014).
71. CMS Collab., JHEP **1307**, 122 (2013).
72. ATLAS Collab., JHEP **1501**, 068 (2015).
73. ATLAS Collab., Phys. Rev. **D92**, 072004 (2015).
74. D0 Collab., Phys. Rev. Lett. **99**, 131801 (2007).
75. ATLAS Collab., Phys. Rev. **D88**, 112003 (2013).
76. CMS Collab., Eur. Phys. J. **C75**, 151 (2015).
77. ATLAS Collab., *Constraints on promptly decaying supersymmetric particles with lepton-number- and R-parity-violating interactions using Run-1 ATLAS data*, ATLAS-CONF-2015-018 (2015).
78. CMS Collab., *Search for RPV SUSY in the four-lepton final state*, CMS-PAS-SUS-13-010 (2013).
79. CDF Collab., Phys. Rev. Lett. **107**, 042001 (2011).
80. ATLAS Collab., Phys. Rev. **D91**, 112016 (2015).
81. CMS Collab., Phys. Lett. **B730**, 193 (2014).
82. ATLAS Collab., Phys. Rev. Lett. **114**, 161801 (2015).
83. H1 Collab., Eur. Phys. J. **C71**, 1572 (2011).
84. CMS Collab., Phys. Rev. **D91**, 052012 (2015).
85. ATLAS Collab., Eur. Phys. J. **C75**, 510 (2015).
86. CDF Collab., Phys. Rev. Lett. **105**, 081802 (2010).
87. D0 Collab., Phys. Lett. **B693**, 95 (2010).
88. CMS Collab., JHEP **1506**, 116 (2015).
89. CMS Collab., JHEP **1303**, 037 (2013), JHEP **1307**, 041 (2013).
90. C. Boehm, A. Djouadi, and Y. Mambrini, Phys. Rev. **D61**, 095006 (2000).
91. CDF Collab., Phys. Rev. **D82**, 092001 (2010).
92. D0 Collab., Phys. Lett. **B696**, 321 (2011).
93. CDF Collab., JHEP **1210**, 158 (2012).
94. D0 Collab., Phys. Lett. **B665**, 1 (2008).
95. CDF Collab., Phys. Rev. Lett. **104**, 251801 (2010).
96. D0 Collab., Phys. Lett. **B674**, 4 (2009).
97. ATLAS Collab., JHEP **1509**, 015 (2014).
98. ATLAS Collab., JHEP **1411**, 118 (2014).
99. CMS Collab., Eur. Phys. J. **C73**, 2667 (2013).



# Searches Particle Listings

## Supersymmetric Particle Searches

100. ZEUS Collab., Eur. Phys. J. **C50**, 269 (2007).
101. CMS Collab., Phys. Rev. Lett. **111**, 221801 (2013).
102. ATLAS Collab., *A search for  $B - L$   $R$ -parity-violating scalar top decays in  $\sqrt{s} = 8$  TeV  $pp$  collisions with the ATLAS experiment*, ATLAS-CONF-2015-015 (2015).
103. CMS Collab., Phys. Rev. Lett. **114**, 061801 (2015).
104. ATLAS Collab., *A search for  $R$ -parity-violating scalar top decays in all-hadronic final states with the ATLAS detector in  $\sqrt{s} = 8$  TeV  $pp$  collisions*, ATLAS-CONF-2015-026 (2015).
105. CMS Collab., Phys. Lett. **B747**, 98 (2015).
106. CDF Collab., Phys. Rev. Lett. **103**, 021802 (2009).
107. D0 Collab., Phys. Rev. **D87**, 052011 (2013).
108. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/01-03.1, <http://lepsusy.web.cern.ch/lepsusy>.
109. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/02-04.1, <http://lepsusy.web.cern.ch/lepsusy>.
110. CDF Collab., *Search for trilepton new physics and chargino-neutralino production at the Collider Detector at Fermilab*, CDF Note 10636 (2011).
111. D0 Collab., Phys. Lett. **B680**, 34 (2009).
112. ATLAS Collab., *Search for the electroweak production of supersymmetric particles in  $\sqrt{s} = 8$  TeV  $pp$  collisions with the ATLAS detector*, arXiv:1509.07152(2015).
113. CMS Collab., Eur. Phys. J. **C74**, 3036 (2014).
114. ATLAS Collab., JHEP **1405**, 071 (2014).
115. ATLAS Collab., Eur. Phys. J. **C75**, 208 (2015).
116. CMS Collab., Phys. Rev. **D90**, 092007 (2014).
117. CMS Collab., *Search for supersymmetry in the vector-boson fusion topology in proton-proton collisions at  $\sqrt{s} = 8$  TeV*, arXiv:1508.07628(2015).
118. ATLAS Collab., Phys. Rev. **D88**, 112006 (2013).
119. CMS Collab., JHEP **1501**, 096 (2015).
120. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/02-05.1, <http://lepsusy.web.cern.ch/lepsusy>.
121. CMS Collab., Eur. Phys. J. **C75**, 325 (2015).
122. LHCb Collab., *Search for long-lived heavy charged particles using a ring imaging Cherenkov technique at LHCb*, arXiv:1506.09173 (2015).
123. H. Dreiner *et al.*, Eur. Phys. J. **C62**, 547 (2009).
124. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/04-07.1, <http://lepsusy.web.cern.ch/lepsusy>.
125. For a sampling of recent post-LHC global analyses, see: P. Bechtle *et al.*, arXiv:1508.05951 (2015); O. Buchmüller *et al.*, Eur. Phys. J. **C74**, 3212 (2014); O. Buchmüller *et al.*, Eur. Phys. J. **C74**, 2922 (2014); M. Citron *et al.*, Phys. Rev. **D87**, 036012 (2013); C. Strey *et al.*, JCAP **1304**, 013 (2013); A. Fowlie *et al.*, Phys. Rev. **D86**, 075010 (2012).
126. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/04-09.1, <http://lepsusy.web.cern.ch/lepsusy>.
127. CDF Collab., Phys. Rev. Lett. **104**, 011801 (2010).
128. D0 Collab., Phys. Rev. Lett. **105**, 221802 (2010).
129. ATLAS Collab., Phys. Rev. **D92**, 072001 (2015).
130. CMS Collab., Phys. Rev. **D92**, 072006 (2015).
131. ATLAS Collab., Eur. Phys. J. **C75**, 318 (2015).
132. CDF Collab., Phys. Rev. **D88**, 031103 (2013).
133. CMS Collab., Phys. Lett. **B722**, 273 (2013).
134. D0 Collab., Phys. Rev. Lett. **101**, 111802 (2008).
135. ATLAS Collab., Phys. Rev. **D90**, 112005 (2014).
136. CMS Collab., *Search for physics beyond the standard model in events with two opposite-sign same-flavor leptons, jets, and missing transverse energy in  $pp$  collisions at  $\sqrt{s} = 8$  TeV*, CMS-PAS-SUS-12-019 (2014).
137. ATLAS Collab., Phys. Rev. **D90**, 052001 (2014).
138. CMS Collab., Phys. Rev. **D91**, 012007 (2015).
139. ATLAS Collab., JHEP **1510**, 134 (2015).
140. CMS Collab., *Phenomenological MSSM interpretation of the CMS 7 and 8 TeV results*, CMS-PAS-SUS-13-020 (2014), CMS-PAS-SUS-15-010 (2015).
141. For a sampling of recent pMSSM analyses, see: K. de Vries *et al.*, Eur. Phys. J. **C75**, 422 (2015); C. Strey *et al.*, JHEP **1409**, 081 (2014); M. Cahill-Rowley *et al.*, Phys. Rev. **D88**, 035002 (2013); C. Boehm *et al.*, JHEP **1306**, 113, (2013); S. AbdusSalam, Phys. Rev. **D87**, 115012 (2013); A. Arbey *et al.*, Eur. Phys. J. **C72**, 2169 (2012); A. Arbey *et al.*, Eur. Phys. J. **C72**, 1847 (2012); M. Carena *et al.*, Phys. Rev. **D86**, 075025 (2012); S. Sekmen *et al.*, JHEP **1202**, 075 (2012).
142. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/04-01.1, <http://lepsusy.web.cern.ch/lepsusy>.
143. ALEPH Collab., Phys. Lett. **B544**, 73 (2002).
144. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/02-09.2, <http://lepsusy.web.cern.ch/lepsusy>.
145. ATLAS Collab., JHEP **1409**, 103 (2014).
146. CMS Collab., Phys. Rev. **D90**, 032006 (2014).
147. CDF Collab., Phys. Rev. Lett. **110**, 201802 (2013).
148. LEP2 SUSY Working Group, ALEPH, DELPHI, L3 and OPAL experiments, note LEPSUSYWG/02-10.1, <http://lepsusy.web.cern.ch/lepsusy>.
149. DELPHI Collab., Eur. Phys. J. **C31**, 412 (2003).
150. T. Falk, K.A. Olive, and M. Srednicki, Phys. Lett. **B339**, 248 (1994).
151. C. Arina and N. Fornengo, JHEP **0711**, 029 (2007).
152. CDF Collab., Phys. Rev. Lett. **105**, 191801 (2010).
153. D0 Collab., Phys. Rev. Lett. **105**, 191802 (2010).
154. ATLAS Collab., Phys. Rev. Lett. **115**, 031801 (2015).
155. CMS Collab., *Search for Lepton Flavour Violating Decays of Heavy Resonances and Quantum Black Holes to electron/muon Pairs in  $pp$  Collisions at a centre of mass energy of 8 TeV*, CMS-PAS-EXO-13-002 (2015).
156. For a sampling of pre-LHC global analyses, see: O. Buchmüller *et al.*, Eur. Phys. J. **C71**, 1722 (2011); E.A. Baltz and P. Gondolo, JHEP **0410**, 052 (2004); B.C. Allanach and C.G. Lester, Phys. Rev. **D73**, 015013 (2006); R.R. de Austri *et al.*, JHEP **0605**, 002 (2006); R. Lafaye *et al.*, Eur. Phys. J. **C54**, 617 (2008); S. Heinemeyer *et al.*, JHEP **0808**, 08 (2008); R. Trotta *et al.*, JHEP **0812**, 024 (2008); P. Bechtle *et al.*, Eur. Phys. J. **C66**, 215 (2010).
157. Supersymmetry Physics Results, ATLAS experiment, <http://twiki.cern.ch/twiki/bin/view/AtlasPublic/SupersymmetryPublicResults/>.



158. Supersymmetry Physics Results, CMS experiment,  
<http://cms-results.web.cern.ch/cms-results/public-results/publications/SUS/index.html>.

### SUPERSYMMETRIC MODEL ASSUMPTIONS

The exclusion of particle masses within a mass range ( $m_1, m_2$ ) will be denoted with the notation “none  $m_1$ – $m_2$ ” in the VALUE column of the following Listings. The latest unpublished results are described in the “Supersymmetry: Experiment” review.

Most of the results shown below, unless stated otherwise, are based on the Minimal Supersymmetric Standard Model (MSSM), as described in the Note on Supersymmetry. Unless otherwise indicated, this includes the assumption of common gaugino and scalar masses at the scale of Grand Unification (GUT), and use of the resulting relations in the spectrum and decay branching ratios. It is also assumed that  $R$ -parity ( $R$ ) is conserved. Unless otherwise indicated, the results also assume that:

- 1) The  $\tilde{\chi}_1^0$  is the lightest supersymmetric particle (LSP)
- 2)  $m_{\tilde{f}_L} = m_{\tilde{f}_R}$ , where  $\tilde{f}_{L,R}$  refer to the scalar partners of left- and right-handed fermions.

Limits involving different assumptions are identified in the Comments or in the Footnotes. We summarize here the notations used in this Chapter to characterize some of the most common deviations from the MSSM (for further details, see the Note on Supersymmetry).

Theories with  $R$ -parity violation ( $\tilde{R}$ ) are characterized by a superpotential of the form:  $\lambda_{ijk} L_i L_j e_k^c + \lambda'_{ijk} L_i Q_j d_k^c + \lambda''_{ijk} u_i^c d_j^c d_k^c$ , where  $i, j, k$  are generation indices. The presence of any of these couplings is often identified in the following by the symbols  $LL\tilde{E}$ ,  $LQ\tilde{D}$ , and  $U\tilde{D}\tilde{D}$ . Mass limits in the presence of  $\tilde{R}$  will often refer to “direct” and “indirect” decays. Direct refers to  $\tilde{R}$  decays of the particle in consideration. Indirect refers to cases where  $\tilde{R}$  appears in the decays of the LSP.

In several models, most notably in theories with so-called Gauge Mediated Supersymmetry Breaking (GMSB), the gravitino ( $\tilde{G}$ ) is the LSP. It is usually much lighter than any other massive particle in the spectrum, and  $m_{\tilde{G}}$  is then neglected in all decay processes involving gravitinos. In these scenarios, particles other than the neutralino are sometimes considered as the next-to-lightest supersymmetric particle (NLSP), and are assumed to decay to their even- $R$  partner plus  $\tilde{G}$ . If the lifetime is short enough for the decay to take place within the detector,  $\tilde{G}$  is assumed to be undetected and to give rise to missing energy ( $\cancel{E}$ ) or missing transverse energy ( $\cancel{E}_T$ ) signatures.

When needed, specific assumptions on the eigenstate content of  $\tilde{\chi}^0$  and  $\tilde{\chi}^\pm$  states are indicated, using the notation  $\tilde{\gamma}$  (photino),  $\tilde{H}$  (higgsino),  $\tilde{W}$  (wino), and  $\tilde{Z}$  (zino) to signal that the limit of pure states was used. The terms gaugino is also used, to generically indicate wino-like charginos and zino-like neutralinos.

### CONTENTS:

- $\tilde{\chi}_1^0$  (Lightest Neutralino) Mass Limit
  - Accelerator limits for stable  $\tilde{\chi}_1^0$
  - Bounds on  $\tilde{\chi}_1^0$  from dark matter searches

- $\tilde{\chi}_1^0$ – $p$  elastic cross section
- Spin-dependent interactions
- Spin-independent interactions
- Other bounds on  $\tilde{\chi}_1^0$  from astrophysics and cosmology
- Unstable  $\tilde{\chi}_1^0$  (Lightest Neutralino) Mass Limit

$\tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$  (Neutralinos) Mass Limits

$\tilde{\chi}_1^\pm, \tilde{\chi}_2^\pm$  (Charginos) Mass Limits

Long-lived  $\tilde{\chi}^\pm$  (Chargino) Mass Limits

$\tilde{\nu}$  (Sneutrino) Mass Limit

Charged Sleptons

–  $\tilde{e}$  (Selectron) Mass Limit

–  $\tilde{\mu}$  (Smuon) Mass Limit

–  $\tilde{\tau}$  (Stau) Mass Limit

– Degenerate Charged Sleptons

–  $\tilde{\ell}$  (Slepton) Mass Limit

$\tilde{q}$  (Squark) Mass Limit

Long-lived  $\tilde{q}$  (Squark) Mass Limit

$\tilde{b}$  (Sbottom) Mass Limit

$\tilde{t}$  (Stop) Mass Limit

Heavy  $\tilde{g}$  (Gluino) Mass Limit

Long-lived/light  $\tilde{g}$  (Gluino) Mass Limit

Light  $\tilde{G}$  (Gravitino) Mass Limits from Collider Experiments

Supersymmetry Miscellaneous Results

### $\tilde{\chi}_1^0$ (Lightest Neutralino) MASS LIMIT

$\tilde{\chi}_1^0$  is often assumed to be the lightest supersymmetric particle (LSP). See also the  $\tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$  section below.

We have divided the  $\tilde{\chi}_1^0$  listings below into five sections:

- 1) Accelerator limits for stable  $\tilde{\chi}_1^0$ ,
- 2) Bounds on  $\tilde{\chi}_1^0$  from dark matter searches,
- 3)  $\tilde{\chi}_1^0$ – $p$  elastic cross section (spin-dependent, spin-independent interactions),
- 4) Other bounds on  $\tilde{\chi}_1^0$  from astrophysics and cosmology, and
- 5) Unstable  $\tilde{\chi}_1^0$  (Lightest Neutralino) mass limit.

### Accelerator limits for stable $\tilde{\chi}_1^0$

Unless otherwise stated, results in this section assume spectra, production rates, decay modes, and branching ratios as evaluated in the MSSM, with gaugino and sfermion mass unification at the GUT scale. These papers generally study production of  $\tilde{\chi}_i^0 \tilde{\chi}_j^0$  ( $i \geq 1, j \geq 2$ ),  $\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ , and (in the case of hadronic collisions)  $\tilde{\chi}_1^\pm \tilde{\chi}_2^0$  pairs. The mass limits on  $\tilde{\chi}_1^0$  are either direct, or follow indirectly from the constraints set by the non-observation of  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_2^0$  states on the gaugino and higgsino MSSM parameters  $M_2$  and  $\mu$ . In some cases, information is used from the nonobservation of slepton decays.

Obsolete limits obtained from  $e^+e^-$  collisions up to  $\sqrt{s}=184$  GeV have been removed from this compilation and can be found in the 2000 Edition (The European Physical Journal **C15** 1 (2000)) of this Review.  $\Delta m = m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ .

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
		1 DREINER 09	THEO	
>40	95	2 ABBIENDI 04H	OPAL	all $\tan\beta$ , $\Delta m > 5$ GeV, $m_0 > 500$ GeV, $A_0 = 0$
>42.4	95	3 HEISTER 04	ALEP	all $\tan\beta$ , all $\Delta m$ , all $m_0$
>39.2	95	4 ABDALLAH 03M	DLPH	all $\tan\beta$ , $m_{\tilde{\nu}} > 500$ GeV
>46	95	5 ABDALLAH 03M	DLPH	all $\tan\beta$ , all $\Delta m$ , all $m_0$
>32.5	95	6 ACCIARRI 00D	L3	$\tan\beta > 0.7$ , $\Delta m > 3$ GeV, all $m_0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		7 AAD 14K	ATLS	

1 DREINER 09 show that in the general MSSM with non-universal gaugino masses there exists no model-independent laboratory bound on the mass of the lightest neutralino. An essentially massless  $\tilde{\chi}_1^0$  is allowed by the experimental and observational data, imposing some constraints on other MSSM parameters, including  $M_2$ ,  $\mu$  and the slepton and squark masses.

2 ABBIENDI 04H search for charginos and neutralinos in events with acoplanar leptons+jets and multi-jet final states in the 192–209 GeV data, combined with the results on leptonic final states from ABBIENDI 04. The results hold for a scan over the parameter space covering the region  $0 < M_2 < 5000$  GeV,  $-1000 < \mu < 1000$  GeV and  $\tan\beta$  from 1 to 40. This limit supersedes ABBIENDI 00H.

3 HEISTER 04 data collected up to 209 GeV. Updates earlier analysis of selectrons from HEISTER 02c, includes a new analysis of charginos and neutralinos decaying into stau and uses results on charginos with initial state radiation from HEISTER 02j. The limit is based on the direct search for charginos and neutralinos, the constraints from the slepton search and the Higgs mass limits from HEISTER 02 using a top mass of 175 GeV, interpreted in a framework with universal gaugino and sfermion masses. Assuming the mixing in the stau sector to be negligible, the limit improves to 43.1 GeV. Under the assumption of MSUGRA with unification of the Higgs and sfermion masses, the limit improves to 50 GeV, and reaches 53 GeV for  $A_0 = 0$ . These limits include and update the results of BARATE 01.

# Searches Particle Listings

## Supersymmetric Particle Searches

<sup>4</sup> ABDALLAH 03M uses data from  $\sqrt{s}=192\text{--}208$  GeV. A limit on the mass of  $\tilde{\chi}_1^0$  is derived from direct searches for neutralinos combined with the chargino search. Neutralinos are searched in the production of  $\tilde{\chi}_1^{0,0}\tilde{\chi}_2^{0,0}$ ,  $\tilde{\chi}_1^{0,0}\tilde{\chi}_3^{0,0}$ , as well as  $\tilde{\chi}_2^{0,0}\tilde{\chi}_3^{0,0}$  and  $\tilde{\chi}_2^{0,0}\tilde{\chi}_4^{0,0}$  giving rise to cascade decays, and  $\tilde{\chi}_1^{0,0}\tilde{\chi}_2^{0,0}$  and  $\tilde{\chi}_1^{0,0}\tilde{\chi}_3^{0,0}$ , followed by the decay  $\tilde{\chi}_2^0 \rightarrow \tilde{\tau}\tau$ . The results hold for the parameter space defined by values of  $M_2 < 1$  TeV,  $|\mu| \leq 2$  TeV with the  $\tilde{\chi}_1^0$  as LSP. The limit is obtained for  $\tan\beta = 1$  and large  $m_0$ , where  $\tilde{\chi}_2^0\tilde{\chi}_4^0$  and chargino pair production are important. If the constraint from Higgs searches is also imposed, the limit improves to 49.0 GeV in the  $m_h^{\text{max}}$  scenario with  $m_t=174.3$  GeV. These limits update the results of ABREU 00J.

<sup>5</sup> ABDALLAH 03M uses data from  $\sqrt{s}=192\text{--}208$  GeV. An indirect limit on the mass of  $\tilde{\chi}_1^0$  is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays and  $\tilde{\tau}\tau$  final states), for charginos (for all  $\Delta m_+$ ) and for sleptons, stop and sbottom. The results hold for the full parameter space defined by values of  $M_2 < 1$  TeV,  $|\mu| \leq 2$  TeV with the  $\tilde{\chi}_1^0$  as LSP. Constraints from the Higgs search in the  $m_h^{\text{max}}$  scenario assuming  $m_t=174.3$  GeV are included. The limit is obtained for  $\tan\beta \geq 5$  when stau mixing leads to mass degeneracy between  $\tilde{\tau}_1$  and  $\tilde{\chi}_1^0$  and the limit is based on  $\tilde{\chi}_2^0$  production followed by its decay to  $\tilde{\tau}_1\tau$ . In the pathological scenario where  $m_0$  and  $|\mu|$  are large, so that the  $\tilde{\chi}_2^0$  production cross section is negligible, and where there is mixing in the stau sector but not in stop nor sbottom, the limit is based on charginos with soft decay products and an ISR photon. The limit then degrades to 39 GeV. See Figs. 40–42 for the dependence of the limit on  $\tan\beta$  and  $m_{\tilde{\nu}}$ . These limits update the results of ABREU 00W.

<sup>6</sup> ACCIARRI 00D data collected at  $\sqrt{s}=189$  GeV. The results hold over the full parameter space defined by  $0.7 \leq \tan\beta \leq 60$ ,  $0 \leq M_2 \leq 2$  TeV,  $m_0 \leq 500$  GeV,  $|\mu| \leq 2$  TeV. The minimum mass limit is reached for  $\tan\beta=1$  and large  $m_0$ . The results of slepton searches from ACCIARRI 99W are used to help set constraints in the region of small  $m_0$ . The limit improves to 48 GeV for  $m_0 \gtrsim 200$  GeV and  $\tan\beta \gtrsim 10$ . See their Figs. 6–8 for the  $\tan\beta$  and  $m_0$  dependence of the limits. Updates ACCIARRI 98F.

<sup>7</sup> AAD 14K sets limits on the  $\chi$ -nucleon spin-dependent and spin-independent cross sections out to  $m_\chi = 10$  TeV.

### Bounds on $\tilde{\chi}_1^0$ from dark matter searches

These papers generally exclude regions in the  $M_2-\mu$  parameter plane assuming that  $\tilde{\chi}_1^0$  is the dominant form of dark matter in the galactic halo. These limits are based on the lack of detection in laboratory experiments, telescopes, or by the absence of a signal in underground neutrino detectors. The latter signal is expected if  $\tilde{\chi}_1^0$  accumulates in the Sun or the Earth and annihilates into high-energy  $\nu$ 's.

VALUE DOCUMENT ID TECN  
• • • We do not use the following data for averages, fits, limits, etc. • • •

1	AARTSEN	15c	ICCB
2	AARTSEN	15E	ICCB
3	ABRAMOWSKI15		HESS
4	ACKERMANN	15	FLAT
5	ACKERMANN	15A	FLAT
6	ACKERMANN	15B	FLAT
7	ADRIAN-MAR.	15	ANTR
8	BUCKLEY	15	THEO
9	CHOI	15	SKAM
10	ALEKSIK	14	MGIC
11	AVRORIN	14	BAIK
12	AARTSEN	13	ICCB
13	AARTSEN	13C	ICCB
14	ABRAMOWSKI13		HESS
15	ADRIAN-MAR.	13	ANTR
16	BERGSTROM	13	COSM
17	BOLIEV	13	BAKS
18	JIN	13	ASTR
16	KOPP	13	COSM
18	ABBASI	12	ICCB
19	ABRAMOWSKI11		HESS
20	ABDO	10	FLAT
21	ACKERMANN	10	FLAT
22	ABBASI	09B	ICCB
23	ACHTERBERG	06	AMND
24	ACKERMANN	06	AMND
25	DEBOER	06	RVUE
26	DESAI	04	SKAM
26	AMBROSIO	99	MCRO
27	LOSECCO	95	RVUE
28	MORI	93	KAMI
29	BOTTINO	92	COSM
30	BOTTINO	91	RVUE
31	GELMINI	91	COSM
32	KAMIONKOW.	91	RVUE
33	MORI	91B	KAMI
34	OLIVE	88	COSM

none 4–15 GeV

<sup>1</sup> AARTSEN 15c is based on 316 live days of running with the IceCube detector. They set a limit of  $1.9 \times 10^{-23} \text{ cm}^3 \text{ s}^{-1}$  on the annihilation cross section to  $\nu\bar{\nu}$  for dark matter with masses between 700–1000 GeV annihilating in the Galactic halo.

<sup>2</sup> AARTSEN 15E is based on 319.7 live days of running with the IceCube 79-string detector. They set a limit of  $4 \times 10^{-24} \text{ cm}^3 \text{ s}^{-1}$  on the annihilation cross section to  $\nu\bar{\nu}$  for dark matter with masses between 30–10000 GeV annihilating in the Galactic center assuming an NFW profile.

<sup>3</sup> ABRAMOWSKI 15 places constraints on the dark matter annihilation cross section for annihilations in the Galactic center for masses between 300 GeV to 10 TeV.

<sup>4</sup> ACKERMANN 15 is based on 5.8 years of data with Fermi-LAT and search for monochromatic gamma-rays in the energy range of 0.2–500 GeV from dark matter annihilations. This updates ACKERMANN 13A.

<sup>5</sup> ACKERMANN 15A is based on 50 months of data with Fermi-LAT and search for dark matter annihilation signals in the isotropic gamma-ray background as well as galactic subhalos in the energy range of a few GeV to a few tens of TeV.

<sup>6</sup> ACKERMANN 15B is based on 6 years of data with Fermi-LAT observations of Milky Way dwarf spheroidal galaxies. Set limits on the annihilation cross section from  $m_\chi = 2$  GeV to 10 TeV. This updates ACKERMANN 14.

<sup>7</sup> ADRIAN-MARTINEZ 15 is based on data from the ANTARES neutrino telescope. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in the galactic center over a background of atmospheric neutrinos and set 90% CL limits on the muon neutrino flux. They also set limits on the annihilation cross section for wimp masses of 25–10000 GeV.

<sup>8</sup> BUCKLEY 15 is based on 5 years of Fermi-LAT data searching for dark matter annihilation signals from Large Magellanic Cloud.

<sup>9</sup> CHOI 15 is based on 3903 days of SuperKamioande data searching for neutrinos produced from dark matter annihilations in the sun. They place constraints on the dark matter-nucleon scattering cross section for dark matter masses between 4–200 GeV.

<sup>10</sup> ALEKSIK 14 is based on almost 160 hours of observations of Segue 1 satellite dwarf galaxy using the MAGIC telescopes between 2011 and 2013. Sets limits on the annihilation cross section out to  $m_\chi = 10$  TeV.

<sup>11</sup> AVRORIN 14 is based on almost 2.76 years with Lake Baikal neutrino telescope. They derive 90% upper limits on the fluxes of muons and muon neutrinos from dark matter annihilations in the Sun.

<sup>12</sup> AARTSEN 13 is based on data collected during 317 effective days with the IceCube 79-string detector including the DeepCore sub-array. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 20–5000 GeV.

<sup>13</sup> AARTSEN 13C is based on data collected during 339.8 effective days with the IceCube 59-string detector. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in nearby galaxies and galaxy clusters. They obtain limits on the neutralino annihilation cross section for neutralino masses in the range 30–100,000 GeV.

<sup>14</sup> ABRAMOWSKI 13 place upper limits on the annihilation cross section with  $\gamma\gamma$  final states in the energy range of 0.5–25 TeV.

<sup>15</sup> ADRIAN-MARTINEZ 13 is based on data from the ANTARES neutrino telescope. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 50–10,000 GeV.

<sup>16</sup> BERGSTROM 13, JIN 13, and KOPP 13 derive limits on the mass and annihilation cross section using AMS-02 data. JIN 13 also sets a limit on the lifetime of the dark matter particle.

<sup>17</sup> BOLIEV 13 is based on data collected during 24.12 years of live time with the Bakson Underground Scintillator Telescope. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 10–1000 GeV.

<sup>18</sup> ABBASI 12 is based on data collected during 812 effective days with AMANDA II and 149 days of the IceCube 40-string detector combined with the data of ABBASI 09B. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. No excess is observed. They also obtain limits on the spin dependent neutralino-proton cross section for neutralino masses in the range 50–5000 GeV.

<sup>19</sup> ABRAMOWSKI 11 place upper limits on the annihilation cross section with  $\gamma\gamma$  final states.

<sup>20</sup> ABDO 10 place upper limits on the annihilation cross section with  $\gamma\gamma$  or  $\mu^+\mu^-$  final states.

<sup>21</sup> ACKERMANN 10 place upper limits on the annihilation cross section with  $b\bar{b}$  or  $\mu^+\mu^-$  final states.

<sup>22</sup> ABBASI 09B is based on data collected during 104.3 effective days with the IceCube 22-string detector. They looked for interactions of  $\nu_\mu$ 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent neutralino-proton cross section for neutralino masses in the range 250–5000 GeV.

<sup>23</sup> ACHTERBERG 06 is based on data collected during 421.9 effective days with the AMANDA detector. They looked for interactions of  $\nu_\mu$ 's from the centre of the Earth over a background of atmospheric neutrinos and set 90 % CL limits on the muon flux. Their limit is compared with the muon flux expected from neutralino annihilations into  $W^+W^-$  and  $b\bar{b}$  at the centre of the Earth for MSSM parameters compatible with the relic dark matter density, see their Fig. 7.

<sup>24</sup> ACKERMANN 06 is based on data collected during 143.7 days with the AMANDA-II detector. They looked for interactions of  $\nu_\mu$ 's from the Sun over a background of atmospheric neutrinos and set 90 % CL limits on the muon flux. Their limit is compared with the muon flux expected from neutralino annihilations into  $W^+W^-$  in the Sun for SUSY model parameters compatible with the relic dark matter density, see their Fig. 3.

<sup>25</sup> DEBOER 06 interpret an excess of diffuse Galactic gamma rays observed with the EGRET satellite as originating from  $\pi^0$  decays from the annihilation of neutralinos into quark jets. They analyze the corresponding parameter space in a supergravity inspired MSSM model with radiative electroweak symmetry breaking, see their Fig. 3 for the preferred region in the  $(m_0, m_{1/2})$  plane of a scenario with large  $\tan\beta$ .

<sup>26</sup> AMBROSIO 99 and DESAI 04 set new neutrino flux limits which can be used to limit the parameter space in supersymmetric models based on neutralino annihilation in the Sun and the Earth.

<sup>27</sup> LOSECCO 95 reanalyzed the IMB data and places lower limit on  $m_{\tilde{\chi}_1^0}$  of 18 GeV if the LSP is a photino and 10 GeV if the LSP is a higgsino based on LSP annihilation in the sun producing high-energy neutrinos and the limits on neutrino fluxes from the IMB detector.

<sup>28</sup> MORI 93 excludes some region in  $M_2-\mu$  parameter space depending on  $\tan\beta$  and lightest scalar Higgs mass for neutralino dark matter  $m_{\tilde{\chi}_0} > m_W$ , using limits on upgoing muons produced by energetic neutrinos from neutralino annihilation in the Sun and the Earth.

- <sup>29</sup> BOTTINO 92 excludes some region  $M_2-\mu$  parameter space assuming that the lightest neutralino is the dark matter, using ongoing muons at Kamiokande, direct searches by Ge detectors, and by LEP experiments. The analysis includes top radiative corrections on Higgs parameters and employs two different hypotheses for nucleon-Higgs coupling. Effects of rescaling in the local neutralino density according to the neutralino relic abundance are taken into account.
- <sup>30</sup> BOTTINO 91 excluded a region in  $M_2-\mu$  plane using ongoing muon data from Kamioka experiment, assuming that the dark matter surrounding us is composed of neutralinos and that the Higgs boson is not too heavy.
- <sup>31</sup> GELMINI 91 exclude a region in  $M_2-\mu$  plane using dark matter searches.
- <sup>32</sup> KAMIONKOWSKI 91 excludes a region in the  $M_2-\mu$  plane using 1MB limit on ongoing muons originated by energetic neutrinos from neutralino annihilation in the sun, assuming that the dark matter is composed of neutralinos and that  $m_{H_1^0} \lesssim 50$  GeV. See Fig. 8 in the paper.
- <sup>33</sup> MORI 91B exclude a part of the region in the  $M_2-\mu$  plane with  $m_{\tilde{\chi}_1^0} \lesssim 80$  GeV using a limit on ongoing muons originated by energetic neutrinos from neutralino annihilation in the earth, assuming that the dark matter surrounding us is composed of neutralinos and that  $m_{H_1^0} \lesssim 80$  GeV.
- <sup>34</sup> OLIVE 88 result assumes that photinos make up the dark matter in the galactic halo. Limit is based on annihilations in the sun and is due to an absence of high energy neutrinos detected in underground experiments. The limit is model dependent.

### $\tilde{\chi}_1^0-p$ elastic cross section

Experimental results on the  $\tilde{\chi}_1^0-p$  elastic cross section are evaluated at  $m_{\tilde{\chi}_1^0}=100$  GeV. The experimental results on the cross section are often mass dependent. Therefore, the mass and cross section results are also given where the limit is strongest, when appropriate. Results are quoted separately for spin-dependent interactions (based on an effective 4-Fermi Lagrangian of the form  $\tilde{\chi}\gamma^\mu\gamma^5\chi\bar{q}\gamma_\mu\gamma^5q$ ) and spin-independent interactions ( $\tilde{\chi}\chi\bar{q}q$ ). For calculational details see GRIEST 88b, ELLIS 88b, BARBIERI 89c, DREES 93b, ARNOWITT 96, BERGSTROM 96, and BAER 97 in addition to the theory papers listed in the Tables. For a description of the theoretical assumptions and experimental techniques underlying most of the listed papers, see the review on "Dark matter" in this "Review of Particle Physics," and references therein. Most of the following papers use galactic halo and nuclear interaction assumptions from (LEWIN 96).

#### Spin-dependent interactions

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.4 \times 10^{-3}$	90	1 AMOLE	15 PICO	C <sub>3</sub> F <sub>8</sub>
$< 6.3 \times 10^{-3}$	90	2 FELIZARDO	14 SMPL	C <sub>2</sub> ClF <sub>5</sub>
$< 0.01$	90	3 APRILE	13 X100	Xe
$< 0.01$	90	4 AKIMOV	12 ZEP3	Xe
$< 0.07$	90	5 ARCHAMBAU..12	PICA	F
$< 7 \times 10^{-3}$		6 BEHNKE	12 COUP	CF <sub>3</sub> I
$< 1.8$	90	7 DAW	12 DRFT	CS <sub>2</sub> ; CF <sub>4</sub>
$< 8.5 \times 10^{-3}$		8 FELIZARDO	12 SMPL	C <sub>2</sub> ClF <sub>5</sub>
$< 0.016$	90	9 KIM	12 KIMS	Csl
$5 \times 10^{-10}$ to $10^{-5}$	95	10 BUCHMUEL...	11B	THEO
$< 1$	90	11 ANGLE	08A	XE10 Xe
$< 0.055$		12 BEDNYAKOV	08	HDMS Ge
$< 0.33$	90	13 BEHNKE	08	COUP CF <sub>3</sub> I
$< 5$		14 AKERIB	06	CDMS Ge
$< 2$		15 SHIMIZU	06A	CNTR CaF <sub>2</sub>
$< 0.4$		16 ALNER	05	NAIA Nal Spin Dep.
$< 2$		17 BARNABE-HE..05	PICA	C
$2 \times 10^{-11}$ to $1 \times 10^{-4}$		18 ELLIS	04	THEO $\mu > 0$
$< 0.8$		19 AHMED	03	NAIA Nal Spin Dep.
$< 40$		20 TAKEDA	03	BOLO NaF Spin Dep.
$< 10$		21 ANGLOHER	02	CRES Sapphire
$8 \times 10^{-7}$ to $2 \times 10^{-5}$		22 ELLIS	01c	THEO $\tan\beta \leq 10$
$< 3.8$		23 BERNABE	00d	DAMA Xe
$< 0.8$		SPoonER	00	UKDM Nal
$< 4.8$		24 BELLI	99c	DAMA F
$< 100$		25 OOTANI	99	BOLO LiF
$< 0.6$		BERNABE	98c	DAMA Xe
$< 5$		24 BERNABE	97	DAMA F

- <sup>1</sup> The strongest limit is 0.001 pb and occurs at  $m_{\tilde{\chi}} = 40$  GeV.
- <sup>2</sup> The strongest limit is 0.0043 pb and occurs at  $m_{\tilde{\chi}} = 35$  GeV. FELIZARDO 14 also presents limits for the scattering on neutrons. At  $m_{\tilde{\chi}} = 100$  GeV, the upper limit is 0.13 pb and the strongest limit is 0.066 pb at  $m_{\tilde{\chi}} = 35$  GeV.
- <sup>3</sup> The strongest limit is 0.006 pb and occurs at  $m_{\tilde{\chi}} = 60$  GeV. APRILE 13 also presents limits for the scattering on neutrons. At 100 GeV, the upper limit is  $4 \times 10^{-4}$  pb and the strongest limit is  $3.5 \times 10^{-4}$  pb at 45 GeV.
- <sup>4</sup> This result updates LEBEDENKO 09A. The strongest limit is  $8 \times 10^{-3}$  pb at  $m_{\tilde{\chi}} = 50$  GeV. Limit applies to the neutralino neutron elastic cross section.
- <sup>5</sup> This result updates ARCHAMBAULT 09. The strongest limit is 0.032 pb at  $m_{\tilde{\chi}} = 20$  GeV.
- <sup>6</sup> The strongest limit is  $6 \times 10^{-3}$  at  $m_{\tilde{\chi}} = 60$  GeV.
- <sup>7</sup> The strongest limit is 1.8 pb and occurs at  $m_{\tilde{\chi}} = 100$  GeV.
- <sup>8</sup> The strongest limit is  $5.7 \times 10^{-3}$  at  $m_{\tilde{\chi}} = 35$  GeV.
- <sup>9</sup> This result updates LEE 07A. The strongest limit is at  $m_{\tilde{\chi}} = 80$  GeV.

- <sup>10</sup> Predictions for the spin-dependent elastic cross section based on a frequentist approach to electroweak observables in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.
- <sup>11</sup> The strongest limit is 0.6 pb and occurs at  $m_{\tilde{\chi}} = 30$  GeV. The limit for scattering on neutrons is 0.01 pb at  $m_{\tilde{\chi}} = 100$  GeV, and the strongest limit is 0.0045 pb at  $m_{\tilde{\chi}} = 30$  GeV.
- <sup>12</sup> Limit applies to neutron elastic cross section.
- <sup>13</sup> The strongest upper limit is 0.25 pb and occurs at  $m_{\tilde{\chi}} \simeq 40$  GeV.
- <sup>14</sup> The strongest upper limit is 4 pb and occurs at  $m_{\tilde{\chi}} \simeq 60$  GeV. The limit on the neutron spin-dependent elastic cross section is 0.07 pb. This latter limit is improved in AHMED 09, where a limit of 0.02 pb is obtained at  $m_{\tilde{\chi}} = 100$  GeV. The strongest limit in AHMED 09 is 0.018 pb and occurs at  $m_{\tilde{\chi}} = 60$  GeV.
- <sup>15</sup> The strongest upper limit is 1.2 pb and occurs at  $m_{\tilde{\chi}} \simeq 40$  GeV. The limit on the neutron spin-dependent cross section is 35 pb.
- <sup>16</sup> The strongest upper limit is 0.35 pb and occurs at  $m_{\tilde{\chi}} \simeq 60$  GeV.
- <sup>17</sup> The strongest upper limit is 1.2 pb and occurs  $m_{\tilde{\chi}} \simeq 30$  GeV.
- <sup>18</sup> ELLIS 04 calculates the  $\chi p$  elastic scattering cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry, but without universal scalar masses. In the case of universal squark and slepton masses, but non-universal Higgs masses, the limit becomes  $2 \times 10^{-4}$ , see ELLIS 03c.
- <sup>19</sup> The strongest upper limit is 0.75 pb and occurs at  $m_{\tilde{\chi}} \approx 70$  GeV.
- <sup>20</sup> The strongest upper limit is 30 pb and occurs at  $m_{\tilde{\chi}} \approx 20$  GeV.
- <sup>21</sup> The strongest upper limit is 8 pb and occurs at  $m_{\tilde{\chi}} \approx 30$  GeV.
- <sup>22</sup> ELLIS 01c calculates the  $\chi-p$  elastic scattering cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry. In models with nonuniversal Higgs masses, the upper limit to the cross section is  $6 \times 10^{-4}$ .
- <sup>23</sup> The strongest upper limit is 3 pb and occurs at  $m_{\tilde{\chi}} \simeq 60$  GeV. The limits are for inelastic scattering  $\chi^0 + {}^{129}\text{Xe} \rightarrow \chi^0 + {}^{129}\text{Xe}^* (39.58 \text{ keV})$ .
- <sup>24</sup> The strongest upper limit is 4.4 pb and occurs at  $m_{\tilde{\chi}} \simeq 60$  GeV.
- <sup>25</sup> The strongest upper limit is about 35 pb and occurs at  $m_{\tilde{\chi}} \simeq 15$  GeV.

#### Spin-independent interactions

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 6.1 \times 10^{-8}$	90	AGNES	15	DSID Ar
	90	1 AGNESE	15A	CDMS Ge
$< 2.2 \times 10^{-8}$	90	2 AGNESE	15B	CDMS Ge
		3 AMOLE	15	PICO C <sub>3</sub> F <sub>8</sub>
$< 1.5 \times 10^{-8}$	90	4 XIAO	15	PANX Xe
		5 AGNESE	14	CDMS Ge
$< 1.5 \times 10^{-9}$	90	6 AKERIB	14	LUX Xe
$10^{-11}-10^{-7}$	95	7 BUCHMUEL...	14A	THEO
$< 4.6 \times 10^{-6}$	90	8 FELIZARDO	14	SMPL C <sub>2</sub> ClF <sub>5</sub>
$10^{-11}-10^{-8}$	95	9 ROSZKOWSKI	14	THEO
		10 AALSETH	13	CGNT Ge
$< 2.2 \times 10^{-6}$	90	11 AGNESE	13	CDMS Si
		12 LI	13b	TEXO Ge
$< 5 \times 10^{-8}$	90	13 AKIMOV	12	ZEP3 Xe
$1.6 \times 10^{-6}; 3.7 \times 10^{-5}$		14 ANGLOHER	12	CRES CaWO <sub>4</sub>
$< 2.6 \times 10^{-9}$	90	15 APRILE	12	X100 Xe
	90	16 ARCHAMBAU..12	PICA	C <sub>4</sub> F <sub>10</sub>
$3 \times 10^{-12}$ to $3 \times 10^{-9}$	95	17 BECHTLE	12	THEO
$< 1.6 \times 10^{-7}$		18 BEHNKE	12	COUP CF <sub>3</sub> I
$< 6.5 \times 10^{-6}$		19 FELIZARDO	12	SMPL C <sub>2</sub> ClF <sub>5</sub>
$< 2.3 \times 10^{-7}$	90	20 KIM	12	KIMS Csl
$< 3.3 \times 10^{-8}$	90	21 AHMED	11A	Ge
$< 4.4 \times 10^{-8}$	90	22 ARMENGAUD	11	EDE2 Ge
$< 7 \times 10^{-7}$	90	23 ANGLOHER	09	CRES CaWO <sub>4</sub>
$< 1 \times 10^{-7}$	90	24 ANGLE	08	XE10 Xe
$< 1 \times 10^{-6}$	90	BENETTI	08	WARP Ar
$< 7.5 \times 10^{-7}$	90	25 ALNER	07A	ZEP2 Xe
$< 2 \times 10^{-7}$		26 AKERIB	06A	CDMS Ge
$< 90 \times 10^{-7}$		ALNER	05	NAIA Nal Spin Indep.
$< 12 \times 10^{-7}$		27 ALNER	05A	ZEPL
$< 20 \times 10^{-7}$		28 ANGLOHER	05	CRES CaWO <sub>4</sub>
$< 14 \times 10^{-7}$		SANGLARD	05	EDEL Ge
$< 4 \times 10^{-7}$		29 AKERIB	04	CDMS Ge
$2 \times 10^{-11}$ to $1.5 \times 10^{-7}$	95	30 BALTZ	04	THEO
$2 \times 10^{-11}$ to $8 \times 10^{-6}$		31,32 ELLIS	04	THEO $\mu > 0$
$< 5 \times 10^{-8}$		33 PIERCE	04A	THEO
$< 2 \times 10^{-5}$		34 AHMED	03	NAIA Nal Spin Indep.
$< 3 \times 10^{-6}$		35 AKERIB	03	CDMS Ge
$2 \times 10^{-13}$ to $2 \times 10^{-7}$		36 BAER	03A	THEO
$< 1.4 \times 10^{-5}$		37 KLAPDOR-K...	03	HDMS Ge
$< 6 \times 10^{-6}$		38 ABRAMS	02	CDMS Ge
$< 1.4 \times 10^{-6}$		39 BENOIT	02	EDEL Ge
$1 \times 10^{-12}$ to $7 \times 10^{-6}$		31 KIM	02b	THEO
$< 3 \times 10^{-5}$		40 MORALES	02b	CSME Ge
$< 1 \times 10^{-5}$		41 MORALES	02c	IGEX Ge
$< 1 \times 10^{-6}$		BALTZ	01	THEO
$< 3 \times 10^{-5}$		42 BAUDIS	01	HDMS Ge
$< 4.5 \times 10^{-6}$		BENOIT	01	EDEL Ge
$< 7 \times 10^{-6}$		43 BOTTINO	01	THEO

# Searches Particle Listings

## Supersymmetric Particle Searches

<  $1 \times 10^{-8}$   
5  $\times 10^{-10}$  to  $1.5 \times 10^{-8}$   
<  $4 \times 10^{-6}$   
2  $\times 10^{-10}$  to  $1 \times 10^{-7}$   
<  $3 \times 10^{-6}$   
<  $6 \times 10^{-7}$   
  
2.5  $\times 10^{-9}$  to  $3.5 \times 10^{-8}$   
<  $1.5 \times 10^{-5}$   
<  $4 \times 10^{-5}$   
<  $7 \times 10^{-6}$   
<  $7 \times 10^{-6}$   
  
44 CORSETTI 01 THEO  $\tan\beta \leq 25$   
45 ELLIS 01c THEO  $\tan\beta \leq 10$   
44 GOMEZ 01 THEO  
44 LAHANAS 01 THEO  
ABUSAIDI 00 CDMS Ge, Si  
46 ACCOMANDO 00 THEO  
47 BERNABEI 00 DAMA Nal  
48 FENG 00 THEO  $\tan\beta=10$   
MORALES 00 IGEX Ge  
SPOONER 00 UKDM Nal  
BAUDIS 99 HDMO  $^{76}\text{Ge}$   
BERNABEI 98c DAMA Xe  
  
1 AGNESE 15A presents 90% CL limits on the elastic cross section for masses in the range 5–20 GeV, from a likelihood analysis of CDMS II data. The limit at 10 GeV is  $2.5 \times 10^{-6}$  pb.  
2 AGNESE 15B result updates AHMED 10 and AHMED 09. The strongest limit is  $1.8 \times 10^{-8}$  pb and occurs at  $m_\chi = 60$  GeV.  
3 AMOLE 15 presents 90% CL limits on the elastic cross section for masses in the range 3–25 GeV. The strongest limit is  $2 \times 10^{-6}$  pb and occurs at  $m_\chi = 25$  GeV.  
4 XIAO 15 presents 90% CL limits on the elastic cross section for masses in the range 3–100 GeV. The strongest limit is  $1 \times 10^{-8}$  pb and occurs at  $m_\chi = 45$  GeV, using the PANDA 54 kg liquid Xenon detector over 80.1 days.  
5 AGNESE 14 presents 90% CL limits on the elastic cross section for masses in the range 3–30 GeV from 577 kg days at SuperCDMS. The strongest limit is  $1 \times 10^{-7}$  pb and occurs at  $m_\chi = 20$  GeV.  
6 The strongest upper limit is  $7.6 \times 10^{-10}$  at  $m_\chi = 33$  GeV.  
7 Predictions for the spin-independent elastic cross section based on a frequentist approach to electroweak observables in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using the 20 fb $^{-1}$  8 TeV and the 5 fb $^{-1}$  7 TeV LHC data and the LUX data.  
8 The strongest limit is  $3.6 \times 10^{-6}$  pb and occurs at  $m_\chi = 35$  GeV.  
9 Predictions for the spin-independent elastic cross section based on a Bayesian approach to electroweak observables in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using the 20 fb $^{-1}$  LHC data and LUX.  
10 AALSETH 13 presents 90% CL limits on the elastic cross section for masses in the range 4–25 GeV in addition to a region of interest at about 8 GeV. The strongest upper limit is  $2 \times 10^{-5}$  pb at  $m_\chi = 14$  GeV.  
11 AGNESE 13 presents 90% CL limits on the elastic cross section for masses in the range 7–100 GeV using the Si based detector. The strongest upper limit is  $1.8 \times 10^{-6}$  pb at  $m_\chi = 50$  GeV. This limit is improved to  $7 \times 10^{-7}$  pb in AGNESE 13A.  
12 LI 13B presents 90% CL limits on the elastic cross section for masses in the range 4–40 GeV. The strongest upper limit is  $4 \times 10^{-5}$  pb at  $m_\chi = 14$  GeV.  
13 This result updates LEBEDENKO 09. The strongest limit is  $3.9 \times 10^{-8}$  pb at  $m_\chi = 52$  GeV.  
14 ANGLÖHER 12 presents results of 730 kg days from the CRESST-II dark matter detector. They find two maxima in the likelihood function corresponding to best fit WIMP masses of 25.3 and 11.6 GeV with elastic cross sections of  $1.6 \times 10^{-6}$  and  $3.7 \times 10^{-5}$  pb respectively, see their Table 4. The statistical significance is more than 4 $\sigma$ .  
15 APRILE 12 updates the result of APRILE 11B. The strongest upper limit is <  $2.0 \times 10^{-9}$  pb and occurs at  $m_\chi \simeq 50$  GeV.  
16 The strongest limit is  $6.1 \times 10^{-5}$  pb at  $m_\chi = 20$  GeV.  
17 Predictions for the spin-independent elastic cross section based on a frequentist approach to electroweak observables in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using the 5 fb $^{-1}$  LHC data and XENON100.  
18 The strongest limit is  $1.4 \times 10^{-7}$  at  $m_\chi = 60$  GeV.  
19 The strongest limit is  $4.7 \times 10^{-6}$  at  $m_\chi = 35$  GeV.  
20 This result updates LEE 07A. The strongest limit is  $2.1 \times 10^{-7}$  at  $m_\chi = 70$  GeV.  
21 AHMED 11A gives combined results from CDMS and EDELWEISS. The strongest limit is at  $m_\chi = 90$  GeV.  
22 ARMENGAUD 11 updates result of ARMENGAUD 10. Strongest limit at  $m_\chi = 85$  GeV.  
23 The strongest upper limit is  $4.8 \times 10^{-7}$  pb and occurs at  $m_\chi = 50$  GeV.  
24 The strongest upper limit is  $5.1 \times 10^{-8}$  pb and occurs at  $m_\chi \simeq 30$  GeV. The values quoted here are based on the analysis performed in ANGLE 08 with the update from SORENSEN 09.  
25 The strongest upper limit is  $6.6 \times 10^{-7}$  pb and occurs at  $m_\chi \simeq 65$  GeV.  
26 AKERIB 06A updates the results of AKERIB 05. The strongest upper limit is  $1.6 \times 10^{-7}$  pb and occurs at  $m_\chi \simeq 60$  GeV.  
27 The strongest upper limit is also close to  $1.0 \times 10^{-6}$  pb and occurs at  $m_\chi \simeq 70$  GeV. BENOIT 06 claim that the discrimination power of ZEPLIN-I measurement (ALNER 05A) is not reliable enough to obtain a limit better than  $1 \times 10^{-3}$  pb. However, SMITH 06 do not agree with the criticisms of BENOIT 06.  
28 The strongest upper limit is also close to  $1.4 \times 10^{-6}$  pb and occurs at  $m_\chi \simeq 70$  GeV.  
29 AKERIB 04 is incompatible with BERNABEI 00 most likely value, under the assumption of standard WIMP-halo interactions. The strongest upper limit is  $4 \times 10^{-7}$  pb and occurs at  $m_\chi \simeq 60$  GeV.  
30 Predictions for the spin-independent elastic cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.  
31 KIM 02 and ELLIS 04 calculate the  $\chi p$  elastic scattering cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry, but without universal scalar masses.  
32 In the case of universal squark and slepton masses, but non-universal Higgs masses, the limit becomes  $2 \times 10^{-6}$  ( $2 \times 10^{-11}$  when constraint from the BNL  $g-2$  experiment are included), see ELLIS 03e. ELLIS 05 display the sensitivity of the elastic scattering cross section to the  $\pi$ -Nucleon  $\Sigma$  term.

33 PIERCE 04A calculates the  $\chi p$  elastic scattering cross section in the framework of models with very heavy scalar masses. See Fig. 2 of the paper.  
34 The strongest upper limit is  $1.8 \times 10^{-5}$  pb and occurs at  $m_\chi \simeq 80$  GeV.  
35 Under the assumption of standard WIMP-halo interactions, Akerib 03 is incompatible with BERNABEI 00 most likely value at the 99.98% CL. See Fig. 4.  
36 BAER 03A calculates the  $\chi p$  elastic scattering cross section in several models including the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.  
37 The strongest upper limit is  $7 \times 10^{-6}$  pb and occurs at  $m_\chi \simeq 30$  GeV.  
38 ABRAMS 02 is incompatible with the DAMA most likely value at the 99.9% CL. The strongest upper limit is  $3 \times 10^{-6}$  pb and occurs at  $m_\chi \simeq 30$  GeV.  
39 BENOIT 02 excludes the central result of DAMA at the 99.8% CL.  
40 The strongest upper limit is  $2 \times 10^{-5}$  pb and occurs at  $m_\chi \simeq 40$  GeV.  
41 The strongest upper limit is  $7 \times 10^{-6}$  pb and occurs at  $m_\chi \simeq 46$  GeV.  
42 The strongest upper limit is  $1.8 \times 10^{-5}$  pb and occurs at  $m_\chi \simeq 32$  GeV  
43 BOTTINO 01 calculates the  $\chi$ - $p$  elastic scattering cross section in the framework of the following supersymmetric models:  $N=1$  supergravity with the radiative breaking of the electroweak gauge symmetry,  $N=1$  supergravity with nonuniversal scalar masses and an effective MSSM model at the electroweak scale.  
44 Calculates the  $\chi$ - $p$  elastic scattering cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.  
45 ELLIS 01c calculates the  $\chi$ - $p$  elastic scattering cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry. ELLIS 02B find a range  $2 \times 10^{-8}$ – $1.5 \times 10^{-7}$  at  $\tan\beta=50$ . In models with nonuniversal Higgs masses, the upper limit to the cross section is  $4 \times 10^{-7}$ .  
46 ACCOMANDO 00 calculate the  $\chi$ - $p$  elastic scattering cross section in the framework of minimal  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry. The limit is relaxed by at least an order of magnitude when models with nonuniversal scalar masses are considered. A subset of the authors in ARNOWITT 02 updated the limit to <  $9 \times 10^{-8}$  ( $\tan\beta < 55$ ).  
47 BERNABEI 00 search for annual modulation of the WIMP signal. The data favor the hypothesis of annual modulation at 4 $\sigma$  and are consistent, for a particular model framework quoted there, with  $m_{\chi^0}=44^{+12}_{-9}$  GeV and a spin-independent  $\chi^0$ -proton cross section of  $(5.4 \pm 1.0) \times 10^{-6}$  pb. See also BERNABEI 01 and BERNABEI 00c.  
48 FENG 00 calculate the  $\chi$ - $p$  elastic scattering cross section in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry with a particular emphasis on focus point models. At  $\tan\beta=50$ , the range is  $8 \times 10^{-8}$ – $4 \times 10^{-7}$ .

### Other bounds on $\tilde{\chi}_1^0$ from astrophysics and cosmology

Most of these papers generally exclude regions in the  $M_2 - \mu$  parameter plane by requiring that the  $\tilde{\chi}_1^0$  contribution to the overall cosmological density is less than some maximal value to avoid overclosure of the Universe. Those not based on the cosmological density are indicated. Many of these papers also include LEP and/or other bounds.

VALUE	DOCUMENT ID	TECN	COMMENT
>46 GeV	1 ELLIS	00	RVUE
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	2 BUCHMUEL...	14	COSM
	3 BUCHMUEL...	14A	COSM
	4 ROSZKOWSKI	14	COSM
	5 CABRERA	13	COSM
	6 ELLIS	13B	COSM
	5 STREGE	13	COSM
	2 AKULA	12	COSM
	2 ARBEY	12A	COSM
	2 BAER	12	COSM
	7 BALAZS	12	COSM
	8 BECHTLE	12	COSM
	9 BESKIDT	12	COSM
> 18 GeV	10 BOTTINO	12	COSM
	2 BUCHMUEL...	12	COSM
	2 CAO	12A	COSM
	2 ELLIS	12B	COSM
	11 FENG	12B	COSM
	2 KADASTIK	12	COSM
	7 STREGE	12	COSM
	12 BUCHMUEL...	11	COSM
	13 ROSZKOWSKI	11	COSM
	14 ELLIS	10	COSM
	15 BUCHMUEL...	09	COSM
	16 DREINER	09	THEO
	17 BUCHMUEL...	08	COSM
	13 ELLIS	08	COSM
	18 CALIBBI	07	COSM
	19 ELLIS	07	COSM
	20 ALLANACH	06	COSM
	21 DE-AUSTRI	06	COSM
	13 BAER	05	COSM
	22 BALTZ	04	COSM
> 6 GeV	10,23 BELANGER	04	THEO
	24 ELLIS	04B	COSM
	25 PIERCE	04A	COSM
	26 BAER	03	COSM
> 6 GeV	10 BOTTINO	03	COSM
	26 CHATTOPAD...	03	COSM

	27	ELLIS	03	COSM	
	13	ELLIS	03B	COSM	
	26	ELLIS	03C	COSM	
	26	LAHANAS	03	COSM	
	28	LAHANAS	02	COSM	
	29	BARGER	01C	COSM	
	30	ELLIS	01B	COSM	
	27	BOEHM	00B	COSM	
	31	FENG	00	COSM	
< 600 GeV	32	ELLIS	98B	COSM	
	33	EDSJO	97	COSM	Co-annihilation
	34	BAER	96	COSM	
	13	BEREZINSKY	95	COSM	
	35	FALK	95	COSM	CP-violating phases
	36	DREES	93	COSM	Minimal supergravity
	37	FALK	93	COSM	Sfermion mixing
	36	KELLEY	93	COSM	Minimal supergravity
	38	MIZUTA	93	COSM	Co-annihilation
	39	LOPEZ	92	COSM	Minimal supergravity, $m_0=A=0$
	40	MCDONALD	92	COSM	
	41	GRIEST	91	COSM	
	42	NOJIRI	91	COSM	Minimal supergravity
	43	OLIVE	91	COSM	
	44	ROSZKOWSKI	91	COSM	
	45	GRIEST	90	COSM	
	43	OLIVE	89	COSM	
none 100 eV – 15 GeV		SREDNICKI	88	COSM	$\tilde{\gamma}; m_{\tilde{t}}=100$ GeV
none 100 eV–5 GeV		ELLIS	84	COSM	$\tilde{\gamma};$ for $m_{\tilde{t}}=100$ GeV
		GOLDBERG	83	COSM	$\tilde{\gamma}$
	46	KRAUSS	83	COSM	$\tilde{\gamma}$
		VYSOTSKI	83	COSM	$\tilde{\gamma}$

<sup>1</sup> ELLIS 00 updates ELLIS 98. Uses LEP  $e^+e^-$  data at  $\sqrt{s}=202$  and 204 GeV to improve bound on neutralino mass to 51 GeV when scalar mass universality is assumed and 46 GeV when Higgs mass universality is relaxed. Limits on  $\tan\beta$  improve to  $> 2.7$  ( $\mu > 0$ ),  $> 2.2$  ( $\mu < 0$ ) when scalar mass universality is assumed and  $> 1.9$  (both signs of  $\mu$ ) when Higgs mass universality is relaxed.

<sup>2</sup> Implications of the LHC result on the Higgs mass and on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.

<sup>3</sup> BUCHMUELLER 14A places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches using the 20  $\text{fb}^{-1}$  8 TeV and the 5  $\text{fb}^{-1}$  7 TeV LHC and the LUX data.

<sup>4</sup> ROSZKOWSKI 14 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using Bayesian statistics and indirect experimental searches using the 20  $\text{fb}^{-1}$  LHC and the LUX data.

<sup>5</sup> CABRERA 13 and STREGE 13 place constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry with and without non-universal Higgs masses using the 5.8  $\text{fb}^{-1}$ ,  $\sqrt{s}=7$  TeV ATLAS supersymmetry searches and XENON100 results.

<sup>6</sup> ELLIS 13B place constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry with and without Higgs mass universality. Models with universality below the GUT scale are also considered.

<sup>7</sup> BALAZS 12 and STREGE 12 place constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using the 1  $\text{fb}^{-1}$  LHC supersymmetry searches, the 5  $\text{fb}^{-1}$  Higgs mass constraints, both with  $\sqrt{s}=7$  TeV, and XENON100 results.

<sup>8</sup> BECHTLE 12 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches, using the 5  $\text{fb}^{-1}$  LHC and XENON100 data.

<sup>9</sup> BESKIDT 12 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches, the 5  $\text{fb}^{-1}$  LHC and the XENON100 data.

<sup>10</sup> BELANGER 04 and BOTTINO 12 (see also BOTTINO 03, BOTTINO 03A and BOTTINO 04) do not assume gaugino or scalar mass unification.

<sup>11</sup> FENG 12B places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry and large sfermion masses using the 1  $\text{fb}^{-1}$  LHC supersymmetry searches, the 5  $\text{fb}^{-1}$  LHC Higgs mass constraints both with  $\sqrt{s}=7$  TeV, and XENON100 results.

<sup>12</sup> BUCHMUELLER 11 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches and including supersymmetry breaking relations between A and B parameters.

<sup>13</sup> Places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry but non-Universal Higgs masses.

<sup>14</sup> ELLIS 10 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry with universality above the GUT scale.

<sup>15</sup> BUCHMUELLER 09 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches.

<sup>16</sup> DREINER 09 show that in the general MSSM with non-universal gaugino masses there exists no model-independent laboratory bound on the mass of the lightest neutralino. An essentially massless  $\tilde{\chi}_1^0$  is allowed by the experimental and observational data, imposing some constraints on other MSSM parameters, including  $M_2$ ,  $\mu$  and the slepton and squark masses.

<sup>17</sup> BUCHMUELLER 08 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches.

<sup>18</sup> CALIBBI 07 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry with universality above the GUT scale including the effects of right-handed neutrinos.

<sup>19</sup> ELLIS 07 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry with universality below the GUT scale.

<sup>20</sup> ALLANACH 06 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.

<sup>21</sup> DE-AUSTRI 06 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.

<sup>22</sup> BALTZ 04 places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.

<sup>23</sup> Limit assumes a pseudo scalar mass  $< 200$  GeV. For larger pseudo scalar masses,  $m_{\tilde{\chi}} > 18(29)$  GeV for  $\tan\beta = 50(10)$ . Bounds from WMAP,  $(g-2)_\mu$ ,  $b \rightarrow s\gamma$ , LEP.

<sup>24</sup> ELLIS 04B places constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry including supersymmetry breaking relations between A and B parameters. See also ELLIS 03D.

<sup>25</sup> PIERCE 04A places constraints on the SUSY parameter space in the framework of models with very heavy scalar masses.

<sup>26</sup> BAER 03, CHATTOPADHYAY 03, ELLIS 03C and LAHANAS 03 place constraints on the SUSY parameter space in the framework of  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry based on WMAP results for the cold dark matter density.

<sup>27</sup> BOEHM 00B and ELLIS 03 place constraints on the SUSY parameter space in the framework of minimal  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry. Includes the effect of  $\tilde{\chi}-\tilde{t}$  co-annihilations.

<sup>28</sup> LAHANAS 02 places constraints on the SUSY parameter space in the framework of minimal  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry. Focuses on the role of pseudo-scalar Higgs exchange.

<sup>29</sup> BARGER 01C use the cosmic relic density inferred from recent CMB measurements to constrain the parameter space in the framework of minimal  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.

<sup>30</sup> ELLIS 01B places constraints on the SUSY parameter space in the framework of minimal  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry. Focuses on models with large  $\tan\beta$ .

<sup>31</sup> FENG 00 explores cosmologically allowed regions of MSSM parameter space with multi-TeV masses.

<sup>32</sup> ELLIS 98B assumes a universal scalar mass and radiative supersymmetry breaking with universal gaugino masses. The upper limit to the LSP mass is increased due to the inclusion of  $\tilde{\chi}-\tilde{t}_R$  coannihilations.

<sup>33</sup> EDSJO 97 included all coannihilation processes between neutralinos and charginos for any neutralino mass and composition.

<sup>34</sup> Notes the location of the neutralino Z resonance and h resonance annihilation corridors in minimal supergravity models with radiative electroweak breaking.

<sup>35</sup> Mass of the bino (=LSP) is limited to  $m_{\tilde{B}} \lesssim 350$  GeV for  $m_t = 174$  GeV.

<sup>36</sup> DREES 93, KELLEY 93 compute the cosmic relic density of the LSP in the framework of minimal  $N=1$  supergravity models with radiative breaking of the electroweak gauge symmetry.

<sup>37</sup> FALK 93 relax the upper limit to the LSP mass by considering sfermion mixing in the MSSM.

<sup>38</sup> MIZUTA 93 include coannihilations to compute the relic density of Higgsino dark matter.

<sup>39</sup> LOPEZ 92 calculate the relic LSP density in a minimal SUSY GUT model.

<sup>40</sup> MCDONALD 92 calculate the relic LSP density in the MSSM including exact tree-level annihilation cross sections for all two-body final states.

<sup>41</sup> GRIEST 91 improve relic density calculations to account for coannihilations, pole effects, and threshold effects.

<sup>42</sup> NOJIRI 91 uses minimal supergravity mass relations between squarks and sleptons to narrow cosmologically allowed parameter space.

<sup>43</sup> Mass of the bino (=LSP) is limited to  $m_{\tilde{B}} \lesssim 350$  GeV for  $m_t \leq 200$  GeV. Mass of the higgsino (=LSP) is limited to  $m_{\tilde{H}} \lesssim 1$  TeV for  $m_t \leq 200$  GeV.

<sup>44</sup> ROSZKOWSKI 91 calculates LSP relic density in mixed gaugino/higgsino region.

<sup>45</sup> Mass of the bino (=LSP) is limited to  $m_{\tilde{B}} \lesssim 550$  GeV. Mass of the higgsino (=LSP) is limited to  $m_{\tilde{H}} \lesssim 3.2$  TeV.

<sup>46</sup> KRAUSS 83 finds  $m_{\tilde{\gamma}}$  not 30 eV to 2.5 GeV. KRAUSS 83 takes into account the gravitino decay. Find that limits depend strongly on reheated temperature. For example a new allowed region  $m_{\tilde{\gamma}} = 4-20$  MeV exists if  $m_{\text{gravitino}} < 40$  TeV. See figure 2.

### Unstable $\tilde{\chi}_1^0$ (Lightest Neutralino) MASS LIMIT

Unless otherwise stated, results in this section assume spectra and production rates as evaluated in the MSSM. Unless otherwise stated, the goldstino or gravitino mass  $m_{\tilde{G}}$  is assumed to be negligible relative to all other masses. In the following,  $\tilde{G}$  is assumed to be undetected and to give rise to a missing energy ( $\cancel{E}$ ) signature.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>380	95	<sup>1</sup> KHACHATRYAN...14L	CMS	$\tilde{\chi}_1^0 \rightarrow Z \tilde{G}$ simplified models,GMSB

# Searches Particle Listings

## Supersymmetric Particle Searches

• • • We do not use the following data for averages, fits, limits, etc. • • •

none 220–380	95	2	AAD	14BH ATLS	$2\gamma + \cancel{E}_T$ , GMSB, SPS8
		3	AAD	13AP ATLS	$2\gamma + \cancel{E}_T$ , GMSB, SPS8
		4	AAD	13Q ATLS	$\gamma + b + \cancel{E}_T$ , higgsino-like neutralino, GMSB
		5	AAD	13R ATLS	$\tilde{\chi}_1^0 \rightarrow \mu j j, R, \lambda'_{211} \neq 0$
>220	95	6	AALTONEN	13i CDF	$\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}, \cancel{E}_T$ , GMSB
		7	CHATRCHYAN	13AH CMS	$\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$ , GMSB, SPS8, $c\tau < 500$ mm
		8	AAD	12CP ATLS	$2\gamma + \cancel{E}_T$ , GMSB
		9	AAD	12CT ATLS	$> 4\ell^\pm, R$
>149	95	10	AAD	12R ATLS	$\tilde{\chi}_1^0 \rightarrow \mu j j, R, \lambda'_{211} \neq 0$
		11	ABAZOV	12AD D0	$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \gamma Z \tilde{G}, \tilde{G}$ , GMSB
		12	CHATRCHYAN	12BK CMS	$2\gamma + \cancel{E}_T$ , GMSB
		13	CHATRCHYAN	11B CMS	$\tilde{W}^0 \rightarrow \gamma \tilde{G}, \tilde{W}^\pm \rightarrow \ell^\pm \tilde{G}$ , GMSB
>175	95	14	AALTONEN	10 CDF	$p\bar{p} \rightarrow \tilde{\chi} \tilde{\chi}, \tilde{\chi} = \tilde{\chi}_2^0, \tilde{\chi}_1^\pm, \tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$ , GMSB
		15	ABAZOV	10P D0	$\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$ , GMSB
		16	ABAZOV	08F D0	$p\bar{p} \rightarrow \tilde{\chi} \tilde{\chi}, \tilde{\chi} = \tilde{\chi}_2^0, \tilde{\chi}_1^\pm, \tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$ , GMSB
		17	ABULENCIA	07H CDF	$R, L\bar{L}\bar{E}$
> 96.8	95	18	ABBIENDI	06B OPAL	$e^+e^- \rightarrow \tilde{B}\tilde{B}, (\tilde{B} \rightarrow \tilde{G}\gamma)$
		19	ABDALLAH	05B DLPH	$e^+e^- \rightarrow \tilde{G}\tilde{\chi}_1^0, (\tilde{\chi}_1^0 \rightarrow \tilde{G}\gamma)$
> 96	95	20	ABDALLAH	05B DLPH	$e^+e^- \rightarrow \tilde{B}\tilde{B}, (\tilde{B} \rightarrow \tilde{G}\gamma)$

<sup>1</sup> KHACHATRYAN 14L searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for evidence of direct pair production of neutralinos with Higgs or  $Z$ -bosons in the decay chain, leading to  $HH, HZ$  and  $ZZ$  final states with missing transverse energy. The decays of 16–20, a Higgs boson to a  $b$ -quark pair, to a photon pair, and to final states with leptons are considered in conjunction with hadronic and leptonic decay modes of the  $Z$  and  $W$  bosons. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of GMSB simplified models where the decays  $\tilde{\chi}_1^0 \rightarrow H\tilde{G}$  or  $\tilde{\chi}_1^0 \rightarrow Z\tilde{G}$  take place either 100% or 50% of the time, see Figs. 16–20.

<sup>2</sup> AAD 14BH searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing non-pointing photons in a diphoton plus missing transverse energy final state. No excess is observed above the background expected from Standard Model processes. The results are used to set 95% C.L. exclusion limits in the context of gauge-mediated supersymmetric breaking models, with the lightest neutralino being the next-to-lightest supersymmetric particle and decaying with a lifetime in the range from 0.25 ns to about 100 ns into a photon and a gravitino. For limits on the NLSF lifetime versus  $\Lambda$  plane, for the SPS8 model, see their Fig. 7.

<sup>3</sup> AAD 13AP searched in 4.8 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing non-pointing photons in a diphoton plus missing transverse energy final state. No excess is observed above the background expected from Standard Model processes. The results are used to set 95% C.L. exclusion limits in the context of gauge-mediated supersymmetric breaking models, with the lightest neutralino being the next-to-lightest supersymmetric particle and decaying with a lifetime in excess of 0.25 ns into a photon and a gravitino. For limits in the NLSF lifetime versus  $\Lambda$  plane, for the SPS8 model, see their Fig. 8.

<sup>4</sup> AAD 13Q searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing a high- $p_T$  isolated photon, at least one jet identified as originating from a bottom quark, and high missing transverse momentum. Such signatures may originate from supersymmetric models with gauge-mediated supersymmetry breaking in events in which one of a pair of higgsino-like neutralinos decays into a photon and a gravitino while the other decays into a Higgs boson and a gravitino. No significant excess above the expected background was found and limits were set on the neutralino mass in a generalized GMSB model (GGM) with a higgsino-like neutralino NLSF, see their Fig. 4. Intermediate neutralino masses between 220 and 380 GeV are excluded at 95% C.L., regardless of the squark and gluino masses, purely on the basis of the expected weak production.

<sup>5</sup> AAD 13R looked in 4.4 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing new, heavy particles that decay at a significant distance from their production point into a final state containing a high-momentum muon and charged hadrons. No excess over the expected background is observed and limits are placed on the production cross-section of neutralinos via squarks for various  $m_{\tilde{q}}, m_{\tilde{\chi}_1^0}$  in an  $R$ -parity violating scenario with

$\lambda'_{211} \neq 0$ , as a function of the neutralino lifetime, see their Fig. 6.

<sup>6</sup> AALTONEN 13i searched in 6.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for events containing  $\cancel{E}_T$  and a delayed photon that arrives late in the detector relative to the time expected from prompt production. No evidence of delayed photon production is observed.

<sup>7</sup> CHATRCHYAN 13AH searched in 4.9 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing  $\cancel{E}_T$  and a delayed photon that arrives late in the detector relative to the time expected from prompt production. No significant excess above the expected background was found and limits were set on the pair production of  $\tilde{\chi}_1^0$  depending on the neutralino proper decay length, see Fig. 8. Supersedes CHATRCHYAN 12BK.

<sup>8</sup> AAD 12CP searched in 4.8 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with two photons and large  $\cancel{E}_T$  due to  $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$  decays in a GMSB framework. No significant excess above the expected background was found and limits were set on the neutralino mass in a generalized GMSB model (GGM) with a bino-like neutralino NLSF, see Figs. 6 and 7. The other sparticle masses were decoupled,  $\tan\beta = 2$  and  $c\tau_{NLSF} < 0.1$  mm. Also, in the framework of the SPS8 model, limits are presented in Fig. 8.

<sup>9</sup> AAD 12CT searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing four or more leptons (electrons or muons) and either moderate values of missing transverse momentum or large effective mass. No significant excess is found in the data. Limits are presented in a simplified model of  $R$ -parity violating supersymmetry in which charginos are pair-produced and then decay into a  $W$ -boson and a  $\tilde{\chi}_1^0$ , which in turn decays through an RPV coupling into two charged leptons ( $e^\pm e^\mp$  or  $\mu^\pm \mu^\mp$ ) and a neutrino. In this model, limits are set on the neutralino mass as a function of the chargino mass, see Fig. 3a. Limits are also set in an  $R$ -parity violating mSUGRA model, see Fig. 3b.

<sup>10</sup> AAD 12R looked in 33 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing new, heavy particles that decay at a significant distance from their production point into a final state containing a high-momentum muon and charged hadrons. No excess over the expected background is observed and limits are placed on the production cross-section

of neutralinos via squarks for various  $(m_{\tilde{q}}, m_{\tilde{\chi}_1^0})$  in an  $R$ -parity violating scenario with

$\lambda'_{211} \neq 0$ , as a function of the neutralino lifetime, see their Fig. 8. Superseded by AAD 13R.

<sup>11</sup> ABAZOV 12AD looked in 6.2 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 1.96$  TeV for events with a photon, a  $Z$ -boson, and large  $\cancel{E}_T$  in the final state. This topology corresponds to a GMSB model where pairs of neutralino NLSPs are either pair produced promptly or from decays of other supersymmetric particles and then decay to either  $Z\tilde{G}$  or  $\gamma\tilde{G}$ . No significant excess over the SM expectation is observed and a limit at 95% C.L. on the cross section is derived as a function of the effective SUSY breaking scale  $\Lambda$ , see Fig. 3. Assuming  $N_{mes} = 2$ ,  $M_{mes} = 3\Lambda$ ,  $\tan\beta = 3$ ,  $\mu = 0.75 M_1$ , and  $C_{grav} = 1$ , the model is excluded at 95% C.L. for values of  $\Lambda < 87$  TeV.

<sup>12</sup> CHATRCHYAN 12BK searched in 2.23 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with two photons and large  $\cancel{E}_T$  due to  $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$  decays in a GMSB framework. No significant excess above the expected background was found and limits were set on the pair production of  $\tilde{\chi}_1^0$  depending on the neutralino lifetime, see Fig. 6.

<sup>13</sup> CHATRCHYAN 11B looked in 35 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with an isolated lepton ( $e$  or  $\mu$ ), a photon and  $\cancel{E}_T$  which may arise in a generalized gauge mediated model from the decay of Wino-like NLSPs. No evidence for an excess over the expected background is observed. Limits are derived in the plane of squark/gluino mass versus Wino mass (see Fig. 4). Mass degeneracy of the produced squarks and gluinos is assumed.

<sup>14</sup> AALTONEN 10 searched in 2.6 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for diphoton events with large  $\cancel{E}_T$ . They may originate from the production of  $\tilde{\chi}^\pm$  in pairs or associated to a  $\tilde{\chi}_1^0$ , decaying into  $\tilde{\chi}_1^0$  which itself decays in GMSB to  $\gamma\tilde{G}$ . There is no excess of events beyond expectation. An upper limit on the cross section is calculated in the GMSB model as a function of the  $\tilde{\chi}_1^0$  mass and lifetime, see their Fig. 2. A limit is derived on the  $\tilde{\chi}_1^0$  mass of 149 GeV for  $\tau_{\tilde{\chi}_1^0} \ll 1$  ns, which improves the results of previous searches.

<sup>15</sup> ABAZOV 10P looked in 6.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for events with at least two isolated  $\gamma$ s and large  $\cancel{E}_T$ . These could be the signature of  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_1^\pm$  production, decaying to  $\tilde{\chi}_1^0$  and finally  $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$  in a GMSB framework. No significant excess over the SM expectation is observed, and a limit at 95% C.L. on the cross section is derived for  $N_{mes} = 1$ ,  $\tan\beta = 15$  and  $\mu > 0$ , see their Fig. 2. This allows them to set a limit on the effective SUSY breaking scale  $\Lambda > 124$  TeV, from which the excluded  $\tilde{\chi}_1^0$  mass range is obtained.

<sup>16</sup> ABAZOV 08F looked in 1.1 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for diphoton events with large  $\cancel{E}_T$ . They may originate from the production of  $\tilde{\chi}^\pm$  in pairs or associated to a  $\tilde{\chi}_1^0$ , decaying to a  $\tilde{\chi}_1^0$  which itself decays promptly in GMSB to  $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$ . No significant excess was found compared to the background expectation. A limit is derived on the masses of SUSY particles in the GMSB framework for  $M = 2\Lambda$ ,  $N = 1$ ,  $\tan\beta = 15$  and  $\mu > 0$ , see Figure 2. It also excludes  $\Lambda < 91.5$  TeV. Supersedes the results of ABAZOV 05A. Superseded by ABAZOV 10P.

<sup>17</sup> ABULENCIA 07H searched in 346 pb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for events with at least three leptons ( $e$  or  $\mu$ ) from the decay of  $\tilde{\chi}_1^0$  via  $L\bar{L}\bar{E}$  couplings. The results are consistent with the hypothesis of no signal. Upper limits on the cross-section are extracted and a limit is derived in the framework of mSUGRA on the masses of  $\tilde{\chi}_1^0$  and  $\tilde{\chi}_1^\pm$ , see e.g. their Fig. 3 and Tab. II.

<sup>18</sup> ABBIENDI 06B use 600 pb<sup>-1</sup> of data from  $\sqrt{s} = 189$ –209 GeV. They look for events with diphotons +  $\cancel{E}$  final states originating from prompt decays of pair-produced neutralinos in a GMSB scenario with  $\tilde{\chi}_1^0$  NLSP. Limits on the cross-section are computed as a function of  $m(\tilde{\chi}_1^0)$ , see their Fig. 14. The limit on the  $\tilde{\chi}_1^0$  mass is for a pure Bino state assuming a prompt decay, with lifetimes up to  $10^{-9}$ s. Supersedes the results of ABBIENDI 04N.

<sup>19</sup> ABDALLAH 05B use data from  $\sqrt{s} = 180$ –209 GeV. They look for events with single photons +  $\cancel{E}$  final states. Limits are computed in the plane  $(m(\tilde{G}), m(\tilde{\chi}_1^0))$ , shown in their Fig. 9b for a pure Bino state in the GMSB framework and in Fig. 9c for a no-scale supergravity model. Supersedes the results of ABREU 00Z.

<sup>20</sup> ABDALLAH 05B use data from  $\sqrt{s} = 130$ –209 GeV. They look for events with diphotons +  $\cancel{E}$  final states and single photons not pointing to the vertex, expected in GMSB when the  $\tilde{\chi}_1^0$  is the NLSP. Limits are computed in the plane  $(m(\tilde{G}), m(\tilde{\chi}_1^0))$ , see their Fig. 10.

The lower limit is derived on the  $\tilde{\chi}_1^0$  mass for a pure Bino state assuming a prompt decay and  $m_{\tilde{e}R} = m_{\tilde{e}L} = 2 m_{\tilde{\chi}_1^0}$ . It improves to 100 GeV for  $m_{\tilde{e}R} = m_{\tilde{e}L} = 1.1 m_{\tilde{\chi}_1^0}$  and

the limit in the plane  $(m(\tilde{\chi}_1^0), m(\tilde{e}R))$  is shown in Fig. 10b. For long-lived neutralinos, cross-section limits are displayed in their Fig. 11. Supersedes the results of ABREU 00Z.

## $\tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$ (Neutralinos) MASS LIMITS

Neutralinos are unknown mixtures of photinos,  $z$ -inos, and neutral higgsinos (the supersymmetric partners of photons and of  $Z$  and Higgs bosons). The limits here apply only to  $\tilde{\chi}_2^0, \tilde{\chi}_3^0$ , and  $\tilde{\chi}_4^0$ .  $\tilde{\chi}_1^0$  is the lightest supersymmetric particle (LSP); see  $\tilde{\chi}_1^0$  Mass Limits. It is not possible to quote rigorous mass limits because they are extremely model dependent; i.e. they depend on branching ratios of various  $\tilde{\chi}^0$  decay modes, on the masses of decay products ( $\tilde{e}, \tilde{\gamma}, \tilde{q}, \tilde{g}$ ), and on the  $\tilde{e}$  mass exchanged in  $e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0$ . Limits arise either from direct searches, or from the MSSM constraints set on the gaugino and higgsino mass parameters  $M_2$  and  $\mu$  through searches for lighter charginos and neutralinos. Often limits are given as contour plots in the  $m_{\tilde{\chi}^0} - m_{\tilde{g}}$  plane vs other parameters. When specific assumptions are made, e.g. the neutralino is a pure photino ( $\tilde{\gamma}$ ), pure  $z$ -ino ( $\tilde{Z}$ ), or pure neutral higgsino ( $\tilde{H}^0$ ), the neutralinos will be labelled as such.

Limits obtained from  $e^+e^-$  collisions at energies up to 136 GeV, as well as other limits from different techniques, are now superseded and have not been included in this compilation. They can be found in the 1998 Edition (The European Physical Journal **C3** 1 (1998)) of this Review. Some later papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>250	95	1 AAD	15BA ATLS	$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
>380	95	2 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow \tau^\pm \nu_{\tilde{\chi}_1^0} \tau^\pm \mp \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\chi}_1^0} = 0$ GeV
>700	95	2 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow \ell^\pm \nu_{\tilde{\chi}_1^0} \ell^\pm \mp \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\chi}_1^0} = 0$ GeV
>345	95	2 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W_{\tilde{\chi}_1^0} Z \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
>148	95	2 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W_{\tilde{\chi}_1^0} H \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
>620	95	3 AAD	14x ATLS	$\geq 4\ell^\pm, \tilde{\chi}_{2,3}^0 \rightarrow \ell^\pm \ell^\mp \tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0$ GeV
		4 AAD	13 ATLS	$3\ell^\pm + \cancel{E}_T$ , pMSSM, SMS
		5 CHATRCHYAN12BJ CMS		$\geq 2\ell, \text{jets} + \cancel{E}_T, pp \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_2^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none	95	6 AAD	14G ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W_{\tilde{\chi}_1^0} Z \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
180–355		7 KHACHATRYAN14I CMS		$\tilde{\chi}_2^0 \rightarrow (Z, H) \tilde{\chi}_1^0 \ell \bar{\ell}$ , simplified model
		8 AAD	12As ATLS	$3\ell^\pm + \cancel{E}_T$ , pMSSM
		9 AAD	12T ATLS	$\ell^\pm \ell^\pm + \cancel{E}_T, pp \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_2^0$

1 AAD 15BA searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of charginos and neutralinos decaying to a final state containing a  $W$  boson and a 125 GeV Higgs boson, plus missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with the decays  $\tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$  and  $\tilde{\chi}_2^0 \rightarrow H \tilde{\chi}_1^0$  having 100% branching fraction, see Fig. 8. A combination of the multiple final states for the Higgs decay yields the best limits (Fig. 8d).

2 AAD 14H searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of charginos and neutralinos decaying to a final state with three leptons and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via either all three generations of leptons, staus only, gauge bosons, or Higgs bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 8.

3 AAD 14x searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the chargino mass in an R-parity conserving simplified model where the decay  $\tilde{\chi}_{2,3}^0 \rightarrow \ell^\pm \ell^\mp \tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 10.

4 AAD 13 searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for charginos and neutralinos decaying to a final state with three leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 and 3, and in simplified models, see Fig. 4. For the simplified models with intermediate slepton decays, degenerate  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_2^0$  masses up to 500 GeV are excluded at 95% C.L. for very large mass differences with the  $\tilde{\chi}_1^0$ . Supersedes AAD 12As.

5 CHATRCHYAN 12BJ searched in 4.98 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for direct electroweak production of charginos and neutralinos in events with at least two leptons, jets and missing transverse momentum. No significant excesses over the expected SM backgrounds are observed and 95% C.L. limits on the production cross section of  $\tilde{\chi}_1^\pm \tilde{\chi}_2^0$  pair production were set in a number of simplified models, see Figs. 7 to 12. Most limits are for exactly 3 jets.

6 AAD 14G searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of chargino-neutralino pairs, decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via gauge bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 10.

7 KHACHATRYAN 14I searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of charginos and neutralinos decaying to a final state with three leptons ( $e$  or  $\mu$ ) and missing transverse momentum, or with a  $Z$ -boson, dijets and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Figs. 12–16.

8 AAD 12As searched in 2.06 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for charginos and neutralinos decaying to a final state with three leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 (top), and in simplified models, see Fig. 2 (bottom).

9 AAD 12T looked in 1 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for the production of supersymmetric particles decaying into final states with missing transverse momentum and exactly two isolated leptons ( $e$  or  $\mu$ ). Same-sign dilepton events were separately studied. Additionally, in opposite-sign events, a search was made for an excess of same-flavor over different-flavor lepton pairs. No excess over the expected background is observed and limits are placed on the effective production cross section of opposite-sign dilepton events with  $\cancel{E}_T > 250$  GeV and on same-sign dilepton events with  $\cancel{E}_T > 100$  GeV. The latter limit is interpreted in a simplified electroweak gaugino production model.

### $\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ (Charginos) MASS LIMITS

Charginos are unknown mixtures of  $w$ -inos and charged higgsinos (the supersymmetric partners of  $W$  and Higgs bosons). A lower mass limit for the lightest chargino ( $\tilde{\chi}_1^\pm$ ) of approximately 45 GeV, independent of the field composition and of the decay mode, has been obtained by the LEP experiments from the analysis of the  $Z$  width and decays. These results, as well as other now superseded limits from  $e^+e^-$  collisions at energies below 136 GeV, and from hadronic collisions, can be found in the 1998 Edition (The European Physical Journal **C3** 1 (1998)) of this Review.

Unless otherwise stated, results in this section assume spectra, production rates, decay modes and branching ratios as evaluated in the MSSM, with gaugino and stfermion mass unification at the GUT scale. These papers generally study production of  $\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ ,  $\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$  and (in the case of hadronic collisions)  $\tilde{\chi}_1^\pm \tilde{\chi}_2^0$  pairs, including the effects of cascade decays. The mass limits on  $\tilde{\chi}_1^\pm$  are either direct, or follow indirectly from the constraints set by the non-observation of  $\tilde{\chi}_2^0$  states on the gaugino and higgsino MSSM parameters  $M_2$  and  $\mu$ . For generic values of the MSSM parameters, limits from high-energy  $e^+e^-$  collisions coincide with the highest value of the mass allowed by phase-space, namely  $m_{\tilde{\chi}_1^\pm} \lesssim \sqrt{s}/2$ . The still unpublished combination of the results of

the four LEP collaborations from the 2000 run of LEP2 at  $\sqrt{s}$  up to  $\approx 209$  GeV yields a lower mass limit of 103.5 GeV valid for general MSSM models. The limits become however weaker in certain regions of the MSSM parameter space where the detection efficiencies or production cross sections are suppressed. For example, this may happen when: (i) the mass differences  $\Delta m_+ = m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0}$  or  $\Delta m_- = m_{\tilde{\chi}_1^\pm} - m_{\tilde{\nu}}$  are very small, and the detection efficiency is reduced; (ii) the electron sneutrino mass is small, and the  $\tilde{\chi}_1^\pm$  production rate is suppressed due to a destructive interference between  $s$  and  $t$  channel exchange diagrams. The regions of MSSM parameter space where the following limits are valid are indicated in the comment lines or in the footnotes.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>250	95	1 AAD	15BA ATLS	$m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
>590	95	2 AAD	15CA ATLS	$\geq 2\gamma + \cancel{E}_T$ , GGM, bino-like NLSP, any NLSP mass
none	95	2 AAD	15CA ATLS	$\geq 1\gamma + e\mu + \cancel{E}_T$ , GGM, wino-like NLSP
124–361				
>700	95	3 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow \ell^\pm \nu_{\tilde{\chi}_1^0} \ell^\pm \mp \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\chi}_1^0} = 0$ GeV
>345	95	3 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W_{\tilde{\chi}_1^0} Z \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
>148	95	3 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W_{\tilde{\chi}_1^0} H \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
>380	95	3 AAD	14H ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow \tau^\pm \nu_{\tilde{\chi}_1^0} \tau^\pm \mp \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}$ , $m_{\tilde{\chi}_1^0} = 0$ GeV
>750	95	4 AAD	14x ATLS	$\geq 4\ell^\pm, \tilde{\chi}_1^\pm \rightarrow W^{(*)\pm} \tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu, \bar{R}$
>210	95	5 KHACHATRYAN14I CMS		$\tilde{\chi}_2^0 \rightarrow H \tilde{\chi}_1^0$ and $\tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$ simplified models, $m_{\tilde{\chi}_2^0} = m_{\tilde{\chi}_1^\pm}$ , $m_{\tilde{\chi}_1^0} = 0$ GeV
		6 AAD	13 ATLS	$3\ell^\pm + \cancel{E}_T$ , pMSSM, SMS
		7 AAD	13B ATLS	$2\ell^\pm + \cancel{E}_T$ , pMSSM, SMS
>540	95	8 AAD	12CT ATLS	$\geq 4\ell^\pm, \cancel{E}_T, m_{\tilde{\chi}_1^0} > 300$ GeV
		9 CHATRCHYAN12BJ CMS		$\geq 2\ell, \text{jets} + \cancel{E}_T, pp \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_2^0$
> 94	95	10 ABDALLAH 03M DLPH		$\tilde{\chi}_1^\pm, \tan\beta \leq 40, \Delta m_+ > 3$ GeV, all $m_0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>410	95	11 AAD	14AV ATLS	$\geq 2\tau + \cancel{E}_T$ , direct $\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ , $\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ production, $m_{\tilde{\chi}_2^0} = m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_1^0} = 0$ GeV
>345	95	12 AAD	14AV ATLS	$\geq 2\tau + \cancel{E}_T$ , direct $\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ production, $m_{\tilde{\chi}_2^0} = 0$ GeV
none	95	13 AAD	14G ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp \rightarrow W^\pm \tilde{\chi}_1^0 W^\mp \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^0} = 0$ GeV
100–105, 120–135, 145–160				
none	95	13 AAD	14G ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp \rightarrow \ell^\pm \nu_{\tilde{\chi}_1^0} \ell^\mp \bar{\nu}_{\tilde{\chi}_1^0}$ , simplified model, $m_{\tilde{\chi}_1^0} = 0$ GeV
none	95	13 AAD	14G ATLS	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0 \rightarrow W_{\tilde{\chi}_1^0} Z \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0} = 0$ GeV
180–355				
>168	95	14 AALTONEN 14 CDF		$3\ell^\pm + \cancel{E}_T, \tilde{\chi}_1^\pm \rightarrow \ell \nu_{\tilde{\chi}_1^0}$ , mSUGRA with $m_0 = 60$ GeV

# Searches Particle Listings

## Supersymmetric Particle Searches

15	KHACHATRYAN 14l	CMS	$\tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \ell\bar{\nu}, \bar{\ell}\nu$ , simplified model
16	AALTONEN 13Q	CDF	$\tilde{\chi}_1^\pm \rightarrow \tau X$ , simplified gravity- and gauge-mediated models
17	AAD 12As	ATLS	$3\ell^\pm + \cancel{E}_T$ , pMSSM
18	AAD 12T	ATLS	$\ell^\pm\ell^\mp + \cancel{E}_T, \ell^\pm\ell^\pm + \cancel{E}_T, p\bar{p} \rightarrow \tilde{\chi}_1^\pm\tilde{\chi}_2^0$
19	CHATRCHYAN 11B	CMS	$\tilde{W}^0 \rightarrow \gamma\tilde{G}, \tilde{W}^\pm \rightarrow \ell^\pm\tilde{G}$ , GMSB
>163	95	20	CHATRCHYAN 11V CMS $\tan\beta=3, m_0=60$ GeV, $A_0=0, \mu>0$

- 1 AAD 15BA searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of charginos and neutralinos decaying to a final state containing a  $W$  boson and a 125 GeV Higgs boson, plus missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with the decays  $\tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$  and  $\tilde{\chi}_2^0 \rightarrow H\tilde{\chi}_1^0$  having 100% branching fraction, see Fig. 8. A combination of the multiple final states for the Higgs decay yields the best limits (Fig. 8d).
- 2 AAD 15CA searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with one or more photons and  $\cancel{E}_T$ , with or without leptons ( $e, \mu$ ). No significant excess above the Standard Model expectation is observed. Limits are set on wino masses in the general gauge-mediated SUSY breaking model (GGM), for wino-like NLSP, see Fig. 9, 12.
- 3 AAD 14H searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of charginos and neutralinos decaying to a final state with three leptons and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via either all three generations of leptons, staus only, gauge bosons, or Higgs bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 8.
- 4 AAD 14X searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectation is observed. Limits are set on the wino-like chargino mass in an  $R$ -parity violating simplified model where the decay  $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\pm\tilde{\chi}_1^0$ , with  $\tilde{\chi}_1^0 \rightarrow \ell^\pm\ell^\mp\nu$ , takes place with a branching ratio of 100%, see Fig. 8.
- 5 KHACHATRYAN 14L searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for evidence of chargino-neutralino  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  pair production with Higgs or  $W$ -bosons in the decay chain, leading to  $HW$  final states with missing transverse energy. The decays of a Higgs boson to a photon pair are considered in conjunction with hadronic and leptonic decay modes of the  $W$  bosons. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of simplified models where the decays  $\tilde{\chi}_2^0 \rightarrow H\tilde{\chi}_1^0$  and  $\tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$  take place 100% of the time, see Figs. 22-23.
- 6 AAD 13 searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for charginos and neutralinos decaying to a final state with three leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 and 3, and in simplified models, see Fig. 4. For the simplified models with intermediate slepton decays, degenerate  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_2^0$  masses up to 500 GeV are excluded at 95% C.L. for very large mass differences with the  $\tilde{\chi}_1^0$ . Supersedes AAD 12As.
- 7 AAD 13B searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for gauginos decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Limits are derived in a simplified model of wino-like chargino pair production, where the chargino always decays to the lightest neutralino via an intermediate on-shell charged slepton, see Fig. 2(b). Chargino masses between 110 and 340 GeV are excluded at 95% C.L. for  $m_{\tilde{\chi}_1^0} = 10$  GeV. Exclusion limits are also derived in the phenomenological MSSM, see Fig. 3.
- 8 AAD 12CT searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing four or more leptons (electrons or muons) and either moderate values of missing transverse momentum or large effective mass. No significant excess is found in the data. Limits are presented in a simplified model of  $R$ -parity violating supersymmetry in which charginos are pair-produced and then decay into a  $W$ -boson and a  $\tilde{\chi}_1^0$ , which in turn decays through an RPV coupling into two charged leptons ( $e^\pm e^\mp$  or  $e^\pm\mu^\mp$ ) and a neutrino. In this model, chargino masses up to 540 GeV are excluded at 95% C.L. for  $m_{\tilde{\chi}_1^0}$  above 300 GeV, see Fig. 3a. The limit deteriorates for lighter  $\tilde{\chi}_1^0$ . Limits are also set in an  $R$ -parity violating mSUGRA model, see Fig. 3b.
- 9 CHATRCHYAN 12BJ searched in 4.98 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for direct electroweak production of charginos and neutralinos in events with at least two leptons, jets and missing transverse momentum. No significant excesses over the expected SM backgrounds are observed and 95% C.L. limits on the production cross section of  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  pair production were set in a number of simplified models, see Figs. 7 to 12.
- 10 ABDALLAH 03M uses data from  $\sqrt{s} = 192$ -208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass of charginos is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays), for charginos and for sleptons. These limits are valid for values of  $M_2 < 1$  TeV,  $|\mu| \leq 2$  TeV with the  $\tilde{\chi}_1^0$  as LSP. Constraints from the Higgs search in the  $m_h^{\max}$  scenario assuming  $m_t = 174.3$  GeV are included. The quoted limit applies if there is no mixing in the third family or when  $m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} > 6$  GeV. If mixing is included the limit degrades to 90 GeV. See Fig. 43 for the mass limits as a function of  $\tan\beta$ . These limits update the results of ABREU 00W.

- 11 AAD 14AV searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for the direct production of charginos, neutralinos and staus in events containing at last two hadronically decaying  $\tau$ -leptons, large missing transverse momentum and low jet activity. The quoted limit was derived for direct  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  and  $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$  production with  $\tilde{\chi}_2^0 \rightarrow \tau\tau \rightarrow \tau\tau\tilde{\chi}_1^0$  and  $\tilde{\chi}_1^\pm \rightarrow \tau\nu(\bar{\nu}\tau) \rightarrow \tau\nu\tilde{\chi}_1^0, m_{\tilde{\chi}_2^0} = m_{\tilde{\chi}_1^\pm}, m_{\tilde{\tau}} = 0.5(m_{\tilde{\chi}_1^\pm} + m_{\tilde{\chi}_1^0}), m_{\tilde{\chi}_1^0} = 0$  GeV. No excess over the expected SM background is observed. Exclusion limits are set in simplified models of  $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$  and  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  pair production, see their Figure 7. Upper limits on the cross section and signal strength for direct di-stau production are derived, see Figures 8 and 9. Also, limits are derived in a pMSSM model where the only light slepton is the  $\tilde{\tau}_R$ , see Figure 10.
- 12 AAD 14AV searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for the direct production of charginos, neutralinos and staus in events containing at last two hadronically decaying  $\tau$ -leptons, large missing transverse momentum and low jet activity. The quoted limit was derived for direct  $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$  production with  $\tilde{\chi}_1^\pm \rightarrow \tau\nu(\bar{\nu}\tau) \rightarrow \tau\nu\tilde{\chi}_1^0, m_{\tilde{\tau}} = 0.5(m_{\tilde{\chi}_1^\pm} + m_{\tilde{\chi}_1^0}), m_{\tilde{\chi}_1^0} = 0$  GeV. No excess over the expected SM background is observed. Exclusion limits are set in simplified models of  $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$  and  $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$  pair production, see their Figure 7. Upper limits on the cross section and signal strength for direct di-stau production are derived, see Figures 8 and 9. Also, limits are derived in a pMSSM model where the only light slepton is the  $\tilde{\tau}_R$ , see Figure 10.
- 13 AAD 14G searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of chargino pairs, or chargino-neutralino pairs, decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of chargino pair production, with chargino decays to the lightest neutralino via either sleptons or gauge bosons, see Fig. 5.; or in simplified models of chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via gauge bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 10.
- 14 AALTONEN 14 searched in 5.8 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for evidence of chargino and next-to-lightest neutralino associated production in final states consisting of three leptons (electrons, muons or taus) and large missing transverse momentum. The results are consistent with the Standard Model predictions within 1.85  $\sigma$ . Limits on the chargino mass are derived in an mSUGRA model with  $m_0 = 60$  GeV,  $\tan\beta = 3$ ,  $A_0 = 0$  and  $\mu > 0$ , see their Fig. 2.
- 15 KHACHATRYAN 14l searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of chargino pairs decaying to a final state with opposite-sign lepton pairs ( $e$  or  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Fig. 18.
- 16 AALTONEN 13Q searched in 6.0 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for evidence of chargino-neutralino associated production in like-sign dilepton final states. One lepton is identified as the hadronic decay of a tau lepton, while the other is an electron or muon. Good agreement with the Standard Model predictions is observed and limits are set on the chargino-neutralino cross section for simplified gravity- and gauge-mediated models, see their Figs. 2 and 3.
- 17 AAD 12AS searched in 2.06 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for charginos and neutralinos decaying to a final state with three leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 (top), and in simplified models, see Fig. 2 (bottom).
- 18 AAD 12T looked in 1 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for the production of supersymmetric particles decaying into final states with missing transverse momentum and exactly two isolated leptons ( $e$  or  $\mu$ ). Opposite-sign and same-sign dilepton events were separately studied. Additionally, in opposite-sign events, a search was made for an excess of same-flavor over different-flavor lepton pairs. No excess over the expected background is observed and limits are placed on the effective production cross section of opposite-sign dilepton events with  $\cancel{E}_T > 250$  GeV and on same-sign dilepton events with  $\cancel{E}_T > 100$  GeV. The latter limit is interpreted in a simplified electroweak gaugino production model as a lower chargino mass limit.
- 19 CHATRCHYAN 11B looked in 35 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=7$  TeV for events with an isolated lepton ( $e$  or  $\mu$ ), a photon and  $\cancel{E}_T$  which may arise in a generalized gauge mediated model from the decay of Wino-like NLSPs. No evidence for an excess over the expected background is observed. Limits are derived in the plane of squark/gluino mass versus Wino mass (see Fig. 4). Mass degeneracy of the produced squarks and gluinos is assumed.
- 20 CHATRCHYAN 11V looked in 35 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with  $\geq 3$  isolated leptons ( $e, \mu$  or  $\tau$ ), with or without jets and  $\cancel{E}_T$ . No evidence for an excess over the expected background is observed. Limits are derived in the CMSSM ( $m_0, m_{1/2}$ ) plane for  $\tan\beta = 3$  (see Fig. 5).

### Long-lived $\tilde{\chi}^\pm$ (Chargino) MASS LIMITS

Limits on charginos which leave the detector before decaying.				
VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>620	95	1 AAD	15AE ATLS	stable $\tilde{\chi}^\pm$
>534	95	2 AAD	15BM ATLS	stable $\tilde{\chi}^\pm$
>239	95	2 AAD	15BM ATLS	$\tilde{\chi}^\pm \rightarrow \tilde{\chi}_1^0\pi^\pm$ , lifetime 1 ns, $m_{\tilde{\chi}^\pm} - m_{\tilde{\chi}_1^0} = 0.14$ GeV
>482	95	2 AAD	15BM ATLS	$\tilde{\chi}^\pm \rightarrow \tilde{\chi}_1^0\pi^\pm$ , lifetime 15 ns, $m_{\tilde{\chi}^\pm} - m_{\tilde{\chi}_1^0} = 0.14$ GeV
>103	95	3 AAD	13H ATLS	long-lived $\tilde{\chi}^\pm \rightarrow \tilde{\chi}_1^0\pi^\pm$ , mAMSB, $\Delta m_{\tilde{\chi}_1^0} = 160$ MeV
> 92	95	4 AAD	12BJ ATLS	long-lived $\tilde{\chi}^\pm \rightarrow \pi^\pm\tilde{\chi}_1^0$ , mAMSB
>171	95	5 ABAZOV	09M D0	$H$
>102	95	6 ABBIENDI	03L OPAL	$\tilde{m}_{\tilde{B}} > 500$ GeV
none 2-93.0	95	7 ABREU	00T DLPH	$\tilde{H}^\pm$ or $m_{\tilde{\nu}} > m_{\tilde{\chi}^\pm}$



See key on page 601

# Searches Particle Listings

## Supersymmetric Particle Searches

• • • We do not use the following data for averages, fits, limits, etc. • • •

>260	95	<sup>8</sup> KHACHATRYAN 15AB CMS	$\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 \pi^\pm, \tau^\pm, \pm 0.2\text{ns}$ , AMSB
>800	95	<sup>9</sup> KHACHATRYAN 15AO CMS	long-lived $\tilde{\chi}_1^\pm$ , mAMSB, $\tau > 100\text{ns}$
>100	95	<sup>9</sup> KHACHATRYAN 15AO CMS	long-lived $\tilde{\chi}_1^\pm$ , mAMSB, $\tau > 3\text{ ns}$
	95	<sup>10</sup> KHACHATRYAN 15W CMS	long-lived $\tilde{\chi}_1^0, \tilde{q} \rightarrow q\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell^+ \ell^- \nu, R$
>270	95	<sup>11</sup> AAD 13BD ATLAS	disappearing-track signature, AMSB
>278	95	<sup>12</sup> ABAZOV 13B D0	long-lived $\tilde{\chi}_1^\pm$ , gaugino-like
>244	95	<sup>12</sup> ABAZOV 13B D0	long-lived $\tilde{\chi}_1^\pm$ , higgsino-like

<sup>1</sup> AAD 15AE searched in  $19.1\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ALTA5 muon system. In the absence of an excess of events above the expected backgrounds, limits are set on stable charginos, see Fig. 10.

<sup>2</sup> AAD 15BM searched in  $18.4\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for stable and metastable non-relativistic charged particles through their anomalous specific ionization energy loss in the ATLAS pixel detector. In absence of an excess of events above the expected backgrounds, limits are set on stable charginos (see Table 5) and on metastable charginos decaying to  $\tilde{\chi}_1^0 \pi^\pm$ , see Fig. 11.

<sup>3</sup> AAD 13H searched in  $4.7\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7\text{ TeV}$  for direct electroweak production of long-lived charginos in the context of AMSB scenarios. The search is based on the signature of a high-momentum isolated track with few associated hits in the outer part of the tracking system, arising from a chargino decay into a neutralino and a low-momentum pion. The  $p_T$  spectrum of the tracks was found to be consistent with the SM expectations. Constraints on the lifetime and the production cross section were obtained, see Fig. 6. In the minimal AMSB framework with  $\tan\beta = 5$ , and  $\mu > 0$ , a chargino having a mass below  $103\text{ (85) GeV}$  for a chargino-neutralino mass splitting  $\Delta m_{\tilde{\chi}_1^\pm}$  of  $160\text{ (170) MeV}$  is excluded at the 95% C.L. See Fig. 7 for more precise bounds.

<sup>4</sup> AAD 12BJ looked in  $1.02\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7\text{ TeV}$  for signatures of decaying charginos resulting in isolated tracks with few associated hits in the outer region of the tracking system. The  $p_T$  spectrum of the tracks was found to be consistent with the SM expectations. Constraints on the lifetime and the production cross section were obtained. In the minimal AMSB framework with  $m_{3/2} < 32\text{ TeV}$ ,  $m_0 < 1.5\text{ TeV}$ ,  $\tan\beta = 5$ , and  $\mu > 0$ , a chargino having a mass below  $92\text{ GeV}$  and a lifetime between  $0.5\text{ ns}$  and  $2\text{ ns}$  is excluded at the 95% C.L. See their Fig. 8 for more precise bounds.

<sup>5</sup> ABAZOV 09M searched in  $1.1\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96\text{ TeV}$  for events with direct production of a pair of charged massive stable particles identified by their TOF. The number of the observed events is consistent with the predicted background. The data are used to constrain the production cross section as a function of the  $\tilde{\chi}_1^\pm$  mass, see their Fig. 2. The quoted limit improves to  $206\text{ GeV}$  for gaugino-like charginos.

<sup>6</sup> ABBENDI 03L used  $e^+e^-$  data at  $\sqrt{s} = 130\text{--}209\text{ GeV}$  to select events with two high momentum tracks with anomalous  $dE/dx$ . The excluded cross section is compared to the theoretical expectation as a function of the heavy particle mass in their Fig. 3. The bounds are valid for colorless fermions with lifetime longer than  $10^{-6}\text{ s}$ . Supersedes the results from ACKERSTAFF 98P.

<sup>7</sup> ABREU 00T searches for the production of heavy stable charged particles, identified by their ionization or Cherenkov radiation, using data from  $\sqrt{s} = 130\text{ to }189\text{ GeV}$ . These limits include and update the results of ABREU 98P.

<sup>8</sup> KHACHATRYAN 15AB searched in  $19.5\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for events containing tracks with little or no associated calorimeter energy deposits and with missing hits in the outer layers of the tracking system (disappearing-track signature). Such disappearing tracks can result from the decay of charginos that are nearly mass degenerate with the lightest neutralino. The number of observed events is in agreement with the background expectation. Limits are set on the cross section of electroweak chargino production in terms of the chargino mass and mean proper lifetime, see Fig. 4. In the minimal AMSB model, a chargino mass below  $260\text{ GeV}$  is excluded at 95% C.L., see their Fig. 5.

<sup>9</sup> KHACHATRYAN 15O searched in  $18.8\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for evidence of long-lived charginos in the context of AMSB and pMSSM scenarios. The results are based on a previously published search for heavy stable charged particles at  $7\text{ and }8\text{ TeV}$ . In the minimal AMSB framework with  $\tan\beta = 5$  and  $\mu \geq 0$ , constraints on the chargino mass and lifetime were placed, see Fig. 5. Charginos with a mass below  $800\text{ (100) GeV}$  are excluded at the 95% C.L. for lifetimes above  $100\text{ ns}$  ( $3\text{ ns}$ ). Constraints are also placed on the pMSSM parameter space, see Fig. 3.

<sup>10</sup> KHACHATRYAN 15W searched in up to  $20.5\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for evidence of long-lived neutralinos produced through  $\tilde{q}\text{-}\bar{q}$  pair production, with  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  and  $\tilde{\chi}_1^0 \rightarrow \ell^+ \ell^- \nu$  ( $R: \lambda_{121}, \lambda_{122} \neq 0$ ). 95% C.L. exclusion limits on cross section times branching ratio are set as a function of mean proper decay length of the neutralino, see Figs. 6 and 9.

<sup>11</sup> AAD 13BD searched in  $20.3\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for events containing tracks with no associated hits in the outer region of the tracking system resulting from the decay of charginos that are nearly mass degenerate with the lightest neutralino, as is often the case in AMSB scenarios. No significant excess above the background expectation is observed for candidate tracks with large transverse momentum. Constraints on chargino properties are obtained and in the minimal AMSB model, a chargino mass below  $270\text{ GeV}$  is excluded at 95% C.L., see their Fig. 7.

<sup>12</sup> ABAZOV 13B looked in  $6.3\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96\text{ TeV}$  for charged massive long-lived particles in events with muon-like particles that have both speed and ionization energy loss inconsistent with muons produced in beam collisions. In the absence of an excess, limits are set at 95% C.L. on gaugino- and higgsino-like charginos, see their Table 20 and Fig. 23.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 400	95	<sup>1</sup> AAD	14X ATLAS	$\geq 4\ell^\pm, \tilde{\nu} \rightarrow \nu\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu, R$
> 94	95	<sup>2</sup> AAD <sup>3</sup> ABDALLAH	11Z ATLAS 03M DLPH	$\tilde{\nu}_\tau \rightarrow e\mu, R$ $1 \leq \tan\beta \leq 40$ , $m_{\tilde{e}_R} - m_{\tilde{\chi}_1^0} > 10\text{ GeV}$
> 84	95	<sup>4</sup> HEISTER	02N ALEP	$\tilde{\nu}_e$ , any $\Delta m$
> 41	95	<sup>5</sup> DECAMP	92 ALEP	$\Gamma(Z \rightarrow \text{invisible}); M(\tilde{\nu})=3$
>2000	95	<sup>6</sup> AAD	15O ATLAS	$\tilde{\nu}_\tau, R (e\mu), \lambda'_{311} = 0.11, \lambda_{33k} = 0.07$
>1700	95	<sup>6</sup> AAD	15O ATLAS	$\tilde{\nu}_\tau, R (\tau\mu, e\tau), \lambda'_{311} = 0.11, \lambda_{33k} = 0.07$
		<sup>7</sup> AAD	13AI ATLAS	$\tilde{\nu}_\tau \rightarrow e\mu, e\tau, \mu\tau, R$
		<sup>8</sup> AAD	11H ATLAS	$\tilde{\nu}_\tau \rightarrow e\mu, R$
		<sup>9</sup> AALTONEN	10Z CDF	$\tilde{\nu}_\tau \rightarrow e\mu, e\tau, \mu\tau, R$
		<sup>10</sup> ABAZOV	10M D0	$\tilde{\nu}_\tau \rightarrow e\mu, R$
> 95	95	<sup>11</sup> ABDALLAH	04H DLPH	AMSB, $\mu > 0$
> 37.1	95	<sup>12</sup> ADRIANI	93M L3	$\Gamma(Z \rightarrow \text{invisible}); M(\tilde{\nu})=1$
> 36	95	ABREU	91F DLPH	$\Gamma(Z \rightarrow \text{invisible}); M(\tilde{\nu})=1$
> 31.2	95	<sup>13</sup> ALEXANDER	91F OPAL	$\Gamma(Z \rightarrow \text{invisible}); M(\tilde{\nu})=1$

<sup>1</sup> AAD 14X searched in  $20.3\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the sneutrino mass in an R-parity violating simplified model where the decay  $\tilde{\nu} \rightarrow \nu\tilde{\chi}_1^0$ , with  $\tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu$ , takes place with a branching ratio of 100%, see Fig. 9.

<sup>2</sup> AAD 11Z looked in  $1.07\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7\text{ TeV}$  for events with one electron and one muon of opposite charge from the production of  $\tilde{\nu}_\tau$  via an  $R \lambda'_{311}$  coupling and followed by a decay via  $\lambda_{312}$  into  $e + \mu$ . No evidence for an  $(e, \mu)$  resonance over the SM expectation is observed, and a limit is derived in the plane of  $\lambda'_{311}$  versus  $m_{\tilde{\nu}}$  for three values of  $\lambda_{312}$ , see their Fig. 2. Masses  $m_{\tilde{\nu}} < 1.32\text{ (1.45) TeV}$  are excluded for  $\lambda'_{311} = 0.10$  and  $\lambda_{312} = 0.05$  ( $\lambda'_{311} = 0.11$  and  $\lambda_{312} = 0.07$ ).

<sup>3</sup> ABDALLAH 03M uses data from  $\sqrt{s} = 192\text{--}208\text{ GeV}$  to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of  $M_2 < 1\text{ TeV}$ ,  $|\mu| \leq 1\text{ TeV}$  with the  $\tilde{\chi}_1^0$  as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of  $\tan\beta$ . These limits update the results of ABREU 00W.

<sup>4</sup> HEISTER 02N derives a bound on  $m_{\tilde{\nu}_e}$  by exploiting the mass relation between the  $\tilde{\nu}_e$  and  $\tilde{e}$ , based on the assumption of universal GUT scale gaugino and scalar masses  $m_{1/2}$  and  $m_0$  and the search described in the  $\tilde{e}$  section. In the MSUGRA framework with radiative electroweak symmetry breaking, the limit improves to  $m_{\tilde{\nu}_e} > 130\text{ GeV}$ , assuming a trilinear coupling  $A_0=0$  at the GUT scale. See Figs. 5 and 7 for the dependence of the limits on  $\tan\beta$ .

<sup>5</sup> DECAMP 92 limit is from  $\Gamma(\text{invisible})/\Gamma(\ell\ell) = 5.91 \pm 0.15$  ( $N_\nu = 2.97 \pm 0.07$ ).

<sup>6</sup> AAD 15O searched in  $20.3\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$  for evidence of heavy particles decaying into  $e\mu, e\tau$  or  $\mu\tau$  final states. No significant excess above the Standard Model expectation is observed, and 95% C.L. exclusions are placed on the cross section times branching ratio for the production of an R-parity-violating supersymmetric tau sneutrino, applicable to any sneutrino flavour, see their Fig. 2.

<sup>7</sup> AAD 13AI searched in  $4.6\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7\text{ TeV}$  for evidence of heavy particles decaying into  $e\mu, e\tau$  or  $\mu\tau$  final states. No significant excess above the Standard Model expectation is observed, and 95% C.L. exclusions are placed on the cross section times branching ratio for the production of an R-parity-violating supersymmetric tau sneutrino, see their Fig. 2. For couplings  $\lambda'_{311} = 0.10$  and  $\lambda_{33k} = 0.05$ , the lower limits on the  $\tilde{\nu}_\tau$  mass are  $1610, 1110, 1100\text{ GeV}$  in the  $e\mu, e\tau$ , and  $\mu\tau$  channels, respectively.

<sup>8</sup> AAD 11H looked in  $35\text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7\text{ TeV}$  for events with one electron and one muon of opposite charge from the production of  $\tilde{\nu}_\tau$  via an  $R \lambda'_{311}$  coupling and followed by a decay via  $\lambda_{312}$  into  $e + \mu$ . No evidence for an excess over the SM expectation is observed, and a limit is derived in the plane of  $\lambda'_{311}$  versus  $m_{\tilde{\nu}}$  for several values of  $\lambda_{312}$ , see their Fig. 2. Superseded by AAD 11Z.

<sup>9</sup> AALTONEN 10Z searched in  $1\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96\text{ TeV}$  for events from the production  $d\bar{d} \rightarrow \tilde{\nu}_\tau$  with the subsequent decays  $\tilde{\nu}_\tau \rightarrow e\mu, \mu\tau, e\tau$  in the MSSM framework with  $R$ . Two isolated leptons of different flavor and opposite charges are required, with  $\tau$ s identified by their hadronic decay. No statistically significant excesses are observed over the SM background. Upper limits on  $\lambda'_{311}$  times the branching ratio are listed in their Table III for various  $\tilde{\nu}_\tau$  masses. Limits on the cross section times branching ratio for  $\lambda'_{311} = 0.10$  and  $\lambda_{33k} = 0.05$ , displayed in Fig. 2, are used to set limits on the  $\tilde{\nu}_\tau$  mass of  $558\text{ GeV}$  for the  $e\mu, 441\text{ GeV}$  for the  $\mu\tau$  and  $442\text{ GeV}$  for the  $e\tau$  channels.

<sup>10</sup> ABAZOV 10M looked in  $5.3\text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96\text{ TeV}$  for events with exactly one pair of high  $p_T$  isolated  $e\mu$  and a veto against hard jets. No evidence for an excess over the SM expectation is observed, and a limit at 95% C.L. on the cross section times branching ratio is derived, see their Fig. 3. These limits are translated into limits on couplings as a function of  $m_{\tilde{\nu}_\tau}$  as shown in their Fig. 4. As an example, for  $m_{\tilde{\nu}_\tau} = 100\text{ GeV}$  and  $\lambda_{312} \leq 0.07$ , couplings  $\lambda'_{311} > 7.7 \times 10^{-4}$  are excluded.

<sup>11</sup> ABDALLAH 04N uses data from LEP 1 and  $\sqrt{s} = 192\text{--}208\text{ GeV}$ . They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region  $1 < m_{3/2} < 50\text{ TeV}$ ,  $0 < m_0 < 1000\text{ GeV}$ ,  $1.5 < \tan\beta < 35$ , both signs of  $\mu$ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z

### $\tilde{\nu}$ (Sneutrino) MASS LIMIT

The limits may depend on the number,  $N(\tilde{\nu})$ , of sneutrinos assumed to be degenerate in mass. Only  $\tilde{\nu}_L$  (not  $\tilde{\nu}_R$ ) is assumed to exist. It is possible that  $\tilde{\nu}$  could be the lightest supersymmetric particle (LSP).

We report here, but do not include in the Listings, the limits obtained from the fit of the final results obtained by the LEP Collaborations on the invisible width of the Z boson ( $\Delta\Gamma_{\text{inv}} < 2.0\text{ MeV}$ , LEP-SLC 06):  $m_{\tilde{\nu}} > 43.7\text{ GeV}$  ( $N(\tilde{\nu})=1$ ) and  $m_{\tilde{\nu}} > 44.7\text{ GeV}$  ( $N(\tilde{\nu})=3$ ).

# Searches Particle Listings

## Supersymmetric Particle Searches

width of 3.2 MeV. The limit is for  $m_t = 174.3$  GeV (see Table 2 for other  $m_t$  values). The limit improves to 114 GeV for  $\mu < 0$ .

<sup>12</sup> ADRIANI 93M limit from  $\Delta\Gamma(Z)(\text{invisible}) < 16.2$  MeV.

<sup>13</sup> ALEXANDER 91F limit is for one species of  $\tilde{\nu}$  and is derived from  $\Gamma(\text{invisible, new})/\Gamma(\ell\ell) < 0.38$ .

### CHARGED SLEPTONS

This section contains limits on charged scalar leptons ( $\tilde{\ell}$ , with  $\ell=e, \mu, \tau$ ). Studies of width and decays of the Z boson (use is made here of  $\Delta\Gamma_{\text{inv}} < 2.0$  MeV, LEP 00) conclusively rule out  $m_{\tilde{\ell}_R} < 40$  GeV (41 GeV for  $\tilde{\ell}_L$ ), independently of decay modes, for each individual slepton. The limits improve to 43 GeV (43.5 GeV for  $\tilde{\ell}_L$ ) assuming all 3 flavors to be degenerate. Limits on higher mass sleptons depend on model assumptions and on the mass splitting  $\Delta m = m_{\tilde{\ell}} - m_{\tilde{\chi}_1^0}$ . The mass and composition

of  $\tilde{\chi}_1^0$  may affect the selection production rate in  $e^+e^-$  collisions through  $t$ -channel exchange diagrams. Production rates are also affected by the potentially large mixing angle of the lightest mass eigenstate  $\tilde{\ell}_1 = \tilde{\ell}_R \sin\theta_{\tilde{\ell}} + \tilde{\ell}_L \cos\theta_{\tilde{\ell}}$ . It is generally assumed that only  $\tilde{\tau}$  may have significant mixing. The coupling to the Z vanishes for  $\theta_{\tilde{\ell}}=0.82$ . In the high-energy limit of  $e^+e^-$  collisions the interference between  $\gamma$  and Z exchange leads to a minimal cross section for  $\theta_{\tilde{\ell}}=0.91$ , a value which is sometimes used in the following entries relative to data taken at LEP2. When limits on  $m_{\tilde{\ell}_R}$  are quoted, it is understood that limits on  $m_{\tilde{\ell}_L}$  are usually at least as strong.

Possibly open decays involving gauginos other than  $\tilde{\chi}_1^0$  will affect the detection efficiencies. Unless otherwise stated, the limits presented here result from the study of  $\tilde{\ell}^+\tilde{\ell}^-$  production, with production rates and decay properties derived from the MSSM. Limits made obsolete by the recent analyses of  $e^+e^-$  collisions at high energies can be found in previous Editions of this Review.

For decays with final state gravitinos ( $\tilde{G}$ ),  $m_{\tilde{G}}$  is assumed to be negligible relative to all other masses.

### $\tilde{e}$ (Selectron) MASS LIMIT

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>410	95	<sup>1</sup> AAD	14x ATLS	$\geq 4\ell^\pm, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu, R$
		<sup>2</sup> CHATRCHYAN14R	CMS	$\geq 3\ell^\pm, \tilde{\ell} \rightarrow \ell^\pm \tau^\mp \tau^\mp \tilde{G}$ simplified model, GMSB, stau (N)NLSP scenario
		<sup>3</sup> AAD	13B ATLS	$2\ell^\pm + \tilde{\ell}_T$ , SMS, pMSSM
> 97.5		<sup>4</sup> ABBIENDI	04 OPAL	$\tilde{e}_R, \Delta m > 11$ GeV, $ \mu  > 100$ GeV, $\tan\beta=1.5$
> 94.4		<sup>5</sup> ACHARD	04 L3	$\tilde{e}_R, \Delta m > 10$ GeV, $ \mu  > 200$ GeV, $\tan\beta \geq 2$
> 71.3		<sup>5</sup> ACHARD	04 L3	$\tilde{e}_R$ , all $\Delta m$
none 30–94	95	<sup>6</sup> ABDALLAH	03M DLPH	$\Delta m > 15$ GeV, $\tilde{e}_R^+ \tilde{e}_R^-$
> 94	95	<sup>7</sup> ABDALLAH	03M DLPH	$\tilde{e}_R, 1 \leq \tan\beta \leq 40, \Delta m > 10$ GeV
> 95	95	<sup>8</sup> HEISTER	02E ALEP	$\Delta m > 15$ GeV, $\tilde{e}_R^+ \tilde{e}_R^-$
> 73	95	<sup>9</sup> HEISTER	02N ALEP	$\tilde{e}_R$ , any $\Delta m$
> 107	95	<sup>9</sup> HEISTER	02N ALEP	$\tilde{e}_L$ , any $\Delta m$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 90–325	95	<sup>10</sup> AAD	14G ATLS	$\tilde{\ell}\tilde{\ell} \rightarrow \ell^+ \tilde{\chi}_1^0 \ell^- \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\ell}_L} = m_{\tilde{\tau}_L}, m_{\tilde{\chi}_1^0} = 0$
		<sup>11</sup> KHACHATRYAN14I	CMS	$\tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$ , simplified model
> 89	95	<sup>12</sup> ABBIENDI	04F OPAL	$R, \tilde{e}_L$
> 92	95	<sup>13</sup> ABDALLAH	04M DLPH	$R, \tilde{e}_R$ , indirect, $\Delta m > 5$ GeV

<sup>1</sup> AAD 14x searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in an R-parity violating simplified model where the decay  $\tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$ , with  $\tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu$ , takes place with a branching ratio of 100%, see Fig. 9.

<sup>2</sup> CHATRCHYAN 14R searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in a stau (N)NLSP simplified model (GMSB) where the decay  $\tilde{\ell} \rightarrow \ell^\pm \tau^\mp \tau^\mp \tilde{G}$  takes place with a branching ratio of 100%, see Fig. 8.

<sup>3</sup> AAD 13b searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for sleptons decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Limits are derived in a simplified model of direct left-handed slepton pair production, where left-handed slepton masses between 85 and 195 GeV are excluded at 95% C.L. for  $m_{\tilde{\chi}_1^0} = 20$  GeV. See also Fig. 2(a). Exclusion

limits are also derived in the phenomenological MSSM, see Fig. 3.

<sup>4</sup> ABBIENDI 04 search for  $\tilde{e}_R \tilde{e}_R$  production in acoplanar di-electron final states in the 183–208 GeV data. See Fig. 13 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$  and for the limit at  $\tan\beta=35$ . This limit supersedes ABBIENDI 00c.

<sup>5</sup> ACHARD 04 search for  $\tilde{e}_R \tilde{e}_L$  and  $\tilde{e}_R \tilde{e}_R$  production in single- and acoplanar di-electron final states in the 192–209 GeV data. Absolute limits on  $m_{\tilde{e}_R}$  are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses

$m_{1/2}$  and  $m_0$ ,  $1 \leq \tan\beta \leq 60$  and  $-2 \leq \mu \leq 2$  TeV. See Fig. 4 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$ . This limit supersedes ACCIARRI 99w.

<sup>6</sup> ABDALLAH 03M looked for acoplanar dielectron +  $\tilde{Z}$  final states at  $\sqrt{s} = 189$ –208 GeV. The limit assumes  $\mu = -200$  GeV and  $\tan\beta=1.5$  in the calculation of the production cross section and  $B(\tilde{e} \rightarrow e\tilde{\chi}_1^0)$ . See Fig. 15 for limits in the  $(m_{\tilde{e}_R}, m_{\tilde{\chi}_1^0})$  plane. These limits

include and update the results of ABREU 01

<sup>7</sup> ABDALLAH 03M uses data from  $\sqrt{s} = 192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of  $M_2 < 1$  TeV,  $|\mu| \leq 1$  TeV with the  $\tilde{\chi}_1^0$  as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of  $\tan\beta$ . These limits update the results of ABREU 00w.

<sup>8</sup> HEISTER 02E looked for acoplanar dielectron +  $\tilde{Z}$  final states from  $e^+e^-$  interactions between 183 and 209 GeV. The mass limit assumes  $\mu < -200$  GeV and  $\tan\beta=2$  for the production cross section and  $B(\tilde{e} \rightarrow e\tilde{\chi}_1^0)=1$ . See their Fig. 4 for the dependence of the limit on  $\Delta m$ . These limits include and update the results of BARATE 01.

<sup>9</sup> HEISTER 02N search for  $\tilde{e}_R \tilde{e}_L$  and  $\tilde{e}_R \tilde{e}_R$  production in single- and acoplanar di-electron final states in the 183–208 GeV data. Absolute limits on  $m_{\tilde{e}_R}$  are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses  $m_{1/2}$  and  $m_0$ ,  $1 \leq \tan\beta \leq 50$  and  $-10 \leq \mu \leq 10$  TeV. The region of small  $|\mu|$ , where cascade decays are important, is covered by a search for  $\tilde{\chi}_1^0 \tilde{\chi}_3^0$  in final states with leptons and possibly photons. Limits on  $m_{\tilde{e}_L}$  are derived by exploiting the mass relation between the  $\tilde{e}_L$  and  $\tilde{e}_R$ , based on universal  $m_0$  and  $m_{1/2}$ . When the constraint from the mass limit of the lightest Higgs from HEISTER 02 is included, the bounds improve to  $m_{\tilde{e}_R} > 77(75)$  GeV and  $m_{\tilde{e}_L} > 115(115)$  GeV for a top mass of 175(180) GeV. In the MSUGRA framework with radiative electroweak symmetry breaking, the limits improve further to  $m_{\tilde{e}_R} > 95$  GeV and  $m_{\tilde{e}_L} > 152$  GeV, assuming a trilinear coupling  $A_0=0$  at the GUT scale. See Figs. 4, 5, 7 for the dependence of the limits on  $\tan\beta$ .

<sup>10</sup> AAD 14G searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of slepton pairs, decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of slepton pair production, see Fig. 8. An interpretation in the pMSSM is also given, see Fig. 10.

<sup>11</sup> KHACHATRYAN 14I searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for electroweak production of slepton pairs decaying to a final state with opposite-sign lepton pairs ( $e$  or  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Fig. 18.

<sup>12</sup> ABBIENDI 04F use data from  $\sqrt{s} = 189$ –209 GeV. They derive limits on sparticle masses under the assumption of  $R$  with  $LL\tilde{E}$  or  $LQ\tilde{D}$  couplings. The results are valid for  $\tan\beta = 1.5$ ,  $\mu = -200$  GeV, with, in addition,  $\Delta m > 5$  GeV for indirect decays via  $LQ\tilde{D}$ . The limit quoted applies to direct decays via  $LL\tilde{E}$  or  $LQ\tilde{D}$  couplings. For indirect decays, the limits on the  $\tilde{e}_R$  mass are respectively 99 and 92 GeV for  $LL\tilde{E}$  and  $LQ\tilde{D}$  couplings and  $m_{\tilde{\chi}_1^0} = 10$  GeV and degrade slightly for larger  $\tilde{\chi}_1^0$  mass. Supersedes the results of ABBIENDI 00.

<sup>13</sup> ABDALLAH 04M use data from  $\sqrt{s} = 192$ –208 GeV to derive limits on sparticle masses under the assumption of  $R$  with  $LL\tilde{E}$  or  $UD\tilde{D}$  couplings. The results are valid for  $\mu = -200$  GeV,  $\tan\beta = 1.5$ ,  $\Delta m > 5$  GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect  $UD\tilde{D}$  decays using the neutralino constraint of 39.5 GeV for  $LL\tilde{E}$  and of 38.0 GeV for  $UD\tilde{D}$  couplings, also derived in ABDALLAH 04M. For indirect decays via  $LL\tilde{E}$  the limit improves to 95 GeV if the constraint from the neutralino is used and to 94 GeV if it is not used. For indirect decays via  $UD\tilde{D}$  couplings it remains unchanged when the neutralino constraint is not used. Supersedes the result of ABREU 00u.

### $\tilde{\mu}$ (Smuon) MASS LIMIT

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>410	95	<sup>1</sup> AAD	14x ATLS	$\geq 4\ell^\pm, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu, R$
		<sup>2</sup> CHATRCHYAN14R	CMS	$\geq 3\ell^\pm, \tilde{\ell} \rightarrow \ell^\pm \tau^\mp \tau^\mp \tilde{G}$ simplified model, GMSB, stau (N)NLSP scenario
		<sup>3</sup> AAD	13B ATLS	$2\ell^\pm + \tilde{\ell}_T$ , SMS, pMSSM
> 91.0		<sup>4</sup> ABBIENDI	04 OPAL	$\Delta m > 3$ GeV, $\tilde{\mu}_R^+ \tilde{\mu}_R^-$ , $ \mu  > 100$ GeV, $\tan\beta=1.5$
> 86.7		<sup>5</sup> ACHARD	04 L3	$\Delta m > 10$ GeV, $\tilde{\mu}_R^+ \tilde{\mu}_R^-$ , $ \mu  > 200$ GeV, $\tan\beta \geq 2$
none 30–88	95	<sup>6</sup> ABDALLAH	03M DLPH	$\Delta m > 5$ GeV, $\tilde{\mu}_R^+ \tilde{\mu}_R^-$
> 94	95	<sup>7</sup> ABDALLAH	03M DLPH	$\tilde{\mu}_R, 1 \leq \tan\beta \leq 40, \Delta m > 10$ GeV
> 98	95	<sup>8</sup> HEISTER	02E ALEP	$\Delta m > 15$ GeV, $\tilde{\mu}_R^+ \tilde{\mu}_R^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
none 90–325	95	<sup>9</sup> AAD	14G ATLS	$\tilde{\ell}\tilde{\ell} \rightarrow \ell^+ \tilde{\chi}_1^0 \ell^- \tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\ell}_L} = m_{\tilde{\tau}_L}, m_{\tilde{\chi}_1^0} = 0$
		<sup>10</sup> KHACHATRYAN14I	CMS	$\tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$ , simplified model
> 87	95	<sup>11</sup> ABDALLAH	04M DLPH	$R, \tilde{\mu}_R$ , indirect, $\Delta m > 5$ GeV
> 81	95	<sup>12</sup> HEISTER	03G ALEP	$\tilde{\mu}_L, R$ decays
> 80	95	<sup>13</sup> ABREU	00V DLPH	$\tilde{\mu}_R \tilde{\mu}_R (\tilde{\mu}_R \rightarrow \mu \tilde{G}), m_{\tilde{G}} > 8$ eV

<sup>1</sup> AAD 14x searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in an R-parity violating simplified model where the decay  $\tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$ , with  $\tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu$ , takes place with a branching ratio of 100%, see Fig. 9.

See key on page 601

## Searches Particle Listings

### Supersymmetric Particle Searches

- <sup>2</sup> CHATRCHYAN 14R searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in a stau (N)NLSP simplified model (GMSB) where the decay  $\tilde{\ell} \rightarrow \ell^\pm \tau^\pm \tilde{\tau}^\mp \tilde{G}$  takes place with a branching ratio of 100%, see Fig. 8.
- <sup>3</sup> AAD 13b searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for sleptons decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Limits are derived in a simplified model of direct left-handed slepton pair production, where left-handed slepton masses between 85 and 195 GeV are excluded at 95% C.L. for  $m_{\tilde{\chi}_1^0} = 20 \text{ GeV}$ . See also Fig. 2(a). Exclusion limits are also derived in the phenomenological MSSM, see Fig. 3.
- <sup>4</sup> ABBIENDI 04 search for  $\tilde{\mu}_R \tilde{\mu}_R$  production in acoplanar di-muon final states in the 183–208 GeV data. See Fig. 14 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$  and for the limit at  $\tan\beta=35$ . Under the assumption of 100% branching ratio for  $\tilde{\mu}_R \rightarrow \mu \tilde{\chi}_1^0$ , the limit improves to 94.0 GeV for  $\Delta m > 4 \text{ GeV}$ . See Fig. 11 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$  at several values of the branching ratio. This limit supersedes ABBIENDI 00g.
- <sup>5</sup> ACHARD 04 search for  $\tilde{\mu}_R \tilde{\mu}_R$  production in acoplanar di-muon final states in the 192–209 GeV data. Limits on  $m_{\tilde{\mu}_R}$  are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses  $m_{1/2}$  and  $m_0$ ,  $1 \leq \tan\beta \leq 60$  and  $-2 \leq \mu \leq 2 \text{ TeV}$ . See Fig. 4 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$ . This limit supersedes ACCIARRI 99w.
- <sup>6</sup> ABDALLAH 03M looked for acoplanar dimuon +  $\cancel{E}_T$  final states at  $\sqrt{s} = 189\text{--}208 \text{ GeV}$ . The limit assumes  $B(\tilde{\mu} \rightarrow \mu \tilde{\chi}_1^0) = 100\%$ . See Fig. 16 for limits on the  $(m_{\tilde{\mu}_R}, m_{\tilde{\chi}_1^0})$  plane. These limits include and update the results of ABREU 01.
- <sup>7</sup> ABDALLAH 03M uses data from  $\sqrt{s} = 192\text{--}208 \text{ GeV}$  to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of  $M_2 < 1 \text{ TeV}$ ,  $|\mu| \leq 1 \text{ TeV}$  with the  $\tilde{\chi}_1^0$  as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of  $\tan\beta$ . These limits update the results of ABREU 00w.
- <sup>8</sup> HEISTER 02E looked for acoplanar dimuon +  $\cancel{E}_T$  final states from  $e^+e^-$  interactions between 183 and 209 GeV. The mass limit assumes  $B(\tilde{\mu} \rightarrow \mu \tilde{\chi}_1^0) = 1$ . See their Fig. 4 for the dependence of the limit on  $\Delta m$ . These limits include and update the results of BARATE 01.
- <sup>9</sup> AAD 14g searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for electroweak production of slepton pairs, decaying to a final state with two leptons ( $e$  and  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of slepton pair production, see Fig. 8. An interpretation in the pMSSM is also given, see Fig. 10.
- <sup>10</sup> KHACHATRYAN 14i searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for electroweak production of slepton pairs decaying to a final state with opposite-sign lepton pairs ( $e$  or  $\mu$ ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Fig. 18.
- <sup>11</sup> ABDALLAH 04M use data from  $\sqrt{s} = 192\text{--}208 \text{ GeV}$  to derive limits on sparticle masses under the assumption of  $R$  with  $L\tilde{L}\tilde{E}$  or  $U\tilde{D}\tilde{D}$  couplings. The results are valid for  $\mu = -200 \text{ GeV}$ ,  $\tan\beta = 1.5$ ,  $\Delta m > 5 \text{ GeV}$  and assuming a BR of 1 for the given decay. The limit quoted is for indirect  $U\tilde{D}\tilde{D}$  decays using the neutralino constraint of 39.5 GeV for  $L\tilde{L}\tilde{E}$  and of 38.0 GeV for  $U\tilde{D}\tilde{D}$  couplings, also derived in ABDALLAH 04M. For indirect decays via  $L\tilde{L}\tilde{E}$  the limit improves to 90 GeV if the constraint from the neutralino is used and remains at 87 GeV if it is not used. For indirect decays via  $U\tilde{D}\tilde{D}$  couplings it degrades to 85 GeV when the neutralino constraint is not used. Supersedes the result of ABREU 00u.
- <sup>12</sup> HEISTER 03G searches for the production of smuons in the case of  $R$  prompt decays with  $L\tilde{L}\tilde{E}$ ,  $L\tilde{Q}\tilde{D}$  or  $U\tilde{D}\tilde{D}$  couplings at  $\sqrt{s} = 189\text{--}209 \text{ GeV}$ . The search is performed for direct and indirect decays, assuming one coupling at a time to be non-zero. The limit holds for direct decays mediated by  $R$   $L\tilde{Q}\tilde{D}$  couplings and improves to 90 GeV for indirect decays (for  $\Delta m > 10 \text{ GeV}$ ). Limits are also given for  $L\tilde{L}\tilde{E}$  direct ( $m_{\tilde{\mu}_R} > 87 \text{ GeV}$ ) and indirect decays ( $m_{\tilde{\mu}_R} > 96 \text{ GeV}$  for  $m(\tilde{\chi}_1^0) > 23 \text{ GeV}$  from BARATE 98s) and for  $U\tilde{D}\tilde{D}$  indirect decays ( $m_{\tilde{\mu}_R} > 85 \text{ GeV}$  for  $\Delta m > 10 \text{ GeV}$ ). Supersedes the results from BARATE 01b.
- <sup>13</sup> ABREU 00v use data from  $\sqrt{s} = 130\text{--}189 \text{ GeV}$  to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of  $m_{\tilde{\tau}_1}$ , after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00q. For limits at different  $m_{\tilde{G}}$ , see their Fig. 12.

#### $\tilde{\tau}$ (Stau) MASS LIMIT

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>85.2		<sup>1</sup> ABBIENDI 04	OPAL	$\Delta m > 6 \text{ GeV}$ , $\theta_\tau = \pi/2$ , $ \mu  > 100 \text{ GeV}$ , $\tan\beta = 1.5$
>78.3		<sup>2</sup> ACHARD 04	L3	$\Delta m > 15 \text{ GeV}$ , $\theta_\tau = \pi/2$ , $ \mu  > 200 \text{ GeV}$ , $\tan\beta \geq 2$
>81.9	95	<sup>3</sup> ABDALLAH 03M	DLPH	$\Delta m > 15 \text{ GeV}$ , all $\theta_\tau$
>79	95	<sup>4</sup> HEISTER 02E	ALEP	$\Delta m > 15 \text{ GeV}$ , $\theta_\tau = \pi/2$
>76	95	<sup>4</sup> HEISTER 02E	ALEP	$\Delta m > 15 \text{ GeV}$ , $\theta_\tau = 0.91$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>5</sup> AAD 12AF	ATLS	$2\tau + \text{jets} + \cancel{E}_T$ , GMSB
		<sup>6</sup> AAD 12AG	ATLS	$\geq 1\tau_h + \text{jets} + \cancel{E}_T$ , GMSB
		<sup>7</sup> AAD 12CM	ATLS	$\geq 1\tau + \text{jets} + \cancel{E}_T$ , GMSB
>87.4	95	<sup>8</sup> ABBIENDI 06B	OPAL	$\tilde{\tau}_R \rightarrow \tau \tilde{G}$ , all $\pi(\tilde{\tau}_R)$
>74	95	<sup>9</sup> ABBIENDI 04F	OPAL	$R$ , $\tilde{\tau}_L$
>68	95	<sup>10</sup> ABDALLAH 04H	DLPH	AMSB, $\mu > 0$
>90	95	<sup>11</sup> ABDALLAH 04M	DLPH	$R$ , $\tilde{\tau}_R$ , indirect, $\Delta m > 5 \text{ GeV}$
none $m_{\tilde{\tau}} - 26.3$	95	<sup>3</sup> ABDALLAH 03M	DLPH	$\Delta m > m_{\tilde{\tau}}$ , all $\theta_\tau$

- <sup>1</sup> ABBIENDI 04 search for  $\tilde{\tau}\tilde{\tau}$  production in acoplanar di-tau final states in the 183–208 GeV data. See Fig. 15 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$  and for the limit at  $\tan\beta=35$ . Under the assumption of 100% branching ratio for  $\tilde{\tau}_R \rightarrow \tau \tilde{\chi}_1^0$ , the limit improves to 89.8 GeV for  $\Delta m > 8 \text{ GeV}$ . See Fig. 12 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$  at several values of the branching ratio and for their dependence on  $\theta_\tau$ . This limit supersedes ABBIENDI 00g.
- <sup>2</sup> ACHARD 04 search for  $\tilde{\tau}\tilde{\tau}$  production in acoplanar di-tau final states in the 192–209 GeV data. Limits on  $m_{\tilde{\tau}_R}$  are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses  $m_{1/2}$  and  $m_0$ ,  $1 \leq \tan\beta \leq 60$  and  $-2 \leq \mu \leq 2 \text{ TeV}$ . See Fig. 4 for the dependence of the limits on  $m_{\tilde{\chi}_1^0}$ .
- <sup>3</sup> ABDALLAH 03M looked for acoplanar ditau +  $\cancel{E}$  final states at  $\sqrt{s} = 130\text{--}208 \text{ GeV}$ . A dedicated search was made for low mass  $\tilde{\tau}$ s decoupling from the  $Z^0$ . The limit assumes  $B(\tilde{\tau} \rightarrow \tau \tilde{\chi}_1^0) = 100\%$ . See Fig. 20 for limits on the  $(m_{\tilde{\tau}}, m_{\tilde{\chi}_1^0})$  plane and as function of the  $\tilde{\chi}_1^0$  mass and of the branching ratio. The limit in the low-mass region improves to 29.6 and 31.1 GeV for  $\tilde{\tau}_R$  and  $\tilde{\tau}_L$ , respectively, at  $\Delta m > m_{\tilde{\tau}}$ . The limit in the high-mass region improves to 84.7 GeV for  $\tilde{\tau}_R$  and  $\Delta m > 15 \text{ GeV}$ . These limits include and update the results of ABREU 01.
- <sup>4</sup> HEISTER 02E looked for acoplanar ditau +  $\cancel{E}_T$  final states from  $e^+e^-$  interactions between 183 and 209 GeV. The mass limit assumes  $B(\tilde{\tau} \rightarrow \tau \tilde{\chi}_1^0) = 1$ . See their Fig. 4 for the dependence of the limit on  $\Delta m$ . These limits include and update the results of BARATE 01.
- <sup>5</sup> AAD 12AF searched in  $2 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with two tau leptons, jets and large  $\cancel{E}_T$  in a GMSB framework. No significant excess above the expected background was found and an upper limit on the visible cross section for new phenomena is set. A 95% C.L. lower limit of 32 TeV on the mGMSB breaking scale  $\Lambda$  is set for  $M_{\text{mess}} = 250 \text{ TeV}$ ,  $N_S = 3$ ,  $\mu > 0$  and  $C_{\text{grav}} = 1$ , independent of  $\tan\beta$ .
- <sup>6</sup> AAD 12AG searched in  $2.05 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with at least one hadronically decaying tau lepton, jets, and large  $\cancel{E}_T$  in a GMSB framework. No significant excess above the expected background was found and an upper limit on the visible cross section for new phenomena is set. A 95% C.L. lower limit of 30 TeV on the mGMSB breaking scale  $\Lambda$  is set for  $M_{\text{mess}} = 250 \text{ TeV}$ ,  $N_S = 3$ ,  $\mu > 0$  and  $C_{\text{grav}} = 1$ , independent of  $\tan\beta$ . For large values of  $\tan\beta$ , the limit on  $\Lambda$  increases to 43 TeV.
- <sup>7</sup> AAD 12CM searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with at least one tau lepton, zero or one additional light lepton ( $e/\mu$ ) jets, and large  $\cancel{E}_T$  in a GMSB framework. No significant excess above the expected background was found and an upper limit on the visible cross section for new phenomena is set. A 95% C. L. lower limit of 54 TeV on the mGMSB breaking scale  $\Lambda$  is set for  $M_{\text{mess}} = 250 \text{ TeV}$ ,  $N_S = 3$ ,  $\mu > 0$  and  $C_{\text{grav}} = 1$ , for  $\tan\beta > 20$ . Here the  $\tilde{\tau}_1$  is the NLSP.
- <sup>8</sup> ABBIENDI 06B use 600 pb $^{-1}$  of data from  $\sqrt{s} = 189\text{--}209 \text{ GeV}$ . They look for events from pair-produced staus in a GMSB scenario with  $\tilde{\tau}$  NLSP including prompt  $\tilde{\tau}$  decays to ditau +  $\cancel{E}$  final states, large impact parameters, kinked tracks and heavy stable charged particles. Limits on the cross-section are computed as a function of  $m(\tilde{\tau})$  and the lifetime, see their Fig. 7. The limit is compared to the  $\sigma \cdot BR^2$  from a scan over the GMSB parameter space.
- <sup>9</sup> ABBIENDI 04F use data from  $\sqrt{s} = 189\text{--}209 \text{ GeV}$ . They derive limits on sparticle masses under the assumption of  $R$  with  $L\tilde{L}\tilde{E}$  or  $L\tilde{Q}\tilde{D}$  couplings. The results are valid for  $\tan\beta = 1.5$ ,  $\mu = -200 \text{ GeV}$ , with, in addition,  $\Delta m > 5 \text{ GeV}$  for indirect decays via  $L\tilde{Q}\tilde{D}$ . The limit quoted applies to direct decays with  $L\tilde{L}\tilde{E}$  couplings and improves to 75 GeV for  $L\tilde{Q}\tilde{D}$  couplings. The limit on the  $\tilde{\tau}_R$  mass for indirect decays is 92 GeV for  $L\tilde{L}\tilde{E}$  couplings at  $m_{\tilde{\chi}_1^0} = 10 \text{ GeV}$  and no exclusion is obtained for  $L\tilde{Q}\tilde{D}$  couplings. Supersedes the results of ABBIENDI 00.
- <sup>10</sup> ABDALLAH 04H use data from LEP 1 and  $\sqrt{s} = 192\text{--}208 \text{ GeV}$ . They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region  $1 < m_{3/2} < 50 \text{ TeV}$ ,  $0 < m_0 < 1000 \text{ GeV}$ ,  $1.5 < \tan\beta < 35$ , both signs of  $\mu$ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM  $Z$  width of 3.2 MeV. The limit is for  $m_{\tilde{t}} = 174.3 \text{ GeV}$  (see Table 2 for other  $m_{\tilde{t}}$  values). The limit improves to 75 GeV for  $\mu < 0$ .
- <sup>11</sup> ABDALLAH 04M use data from  $\sqrt{s} = 192\text{--}208 \text{ GeV}$  to derive limits on sparticle masses under the assumption of  $R$  with  $L\tilde{L}\tilde{E}$  couplings. The results are valid for  $\mu = -200 \text{ GeV}$ ,  $\tan\beta = 1.5$ ,  $\Delta m > 5 \text{ GeV}$  and assuming a BR of 1 for the given decay. The limit quoted is for indirect decays using the neutralino constraint of 39.5 GeV, also derived in ABDALLAH 04M. For indirect decays via  $L\tilde{L}\tilde{E}$  the limit decreases to 86 GeV if the constraint from the neutralino is not used. Supersedes the result of ABREU 00u.

#### Degenerate Charged Sleptons

Unless stated otherwise in the comment lines or in the footnotes, the following limits assume 3 families of degenerate charged sleptons.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>93	95	<sup>1</sup> BARATE 01	ALEP	$\Delta m > 10 \text{ GeV}$ , $\tilde{\ell}_R^+ \tilde{\ell}_R^-$
>70	95	<sup>1</sup> BARATE 01	ALEP	all $\Delta m$ , $\tilde{\ell}_R^+ \tilde{\ell}_R^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>91.9	95	<sup>2</sup> ABBIENDI 06B	OPAL	$\tilde{\ell}_R \rightarrow \ell \tilde{G}$ , all $\ell(\tilde{\ell}_R)$
>88		<sup>3</sup> ABDALLAH 03D	DLPH	$\tilde{\ell}_R \rightarrow \ell \tilde{G}$ , all $\ell(\tilde{\ell}_R)$
>82.7	95	<sup>4</sup> ACHARD 02	L3	$\tilde{\ell}_R$ , $R$ decays, MSUGRA
>83	95	<sup>5</sup> ABBIENDI 01	OPAL	$e^+e^- \rightarrow \tilde{\ell}_1 \tilde{\ell}_1$ , GMSB, $\tan\beta=2$
		<sup>6</sup> ABREU 01	DLPH	$\tilde{\ell} \rightarrow \ell \tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow \gamma \tilde{\chi}_1^0$ , $\ell=e,\mu$
>68.8	95	<sup>7</sup> ACCIARRI 01	L3	$\tilde{\ell}_R, R$ , $0.7 \leq \tan\beta \leq 40$
>84	95	<sup>8</sup> ABREU 00v	DLPH	$\tilde{\ell}_R \tilde{\ell}_R (\tilde{\ell}_R \rightarrow \ell \tilde{G})$ , $m_{\tilde{G}} > 9 \text{ eV}$

# Searches Particle Listings

## Supersymmetric Particle Searches

- <sup>1</sup> BARATE 01 looked for acoplanar dilepton +  $\cancel{E}_T$  and single electron (for  $\tilde{e}\tilde{e}_L$ ) final states at 189 to 202 GeV. The limit assumes  $\mu=200$  GeV and  $\tan\beta=2$  for the production cross section and decay branching ratios, evaluated within the MSSM, and zero efficiency for decays other than  $\tilde{e} \rightarrow \ell\tilde{\chi}_1^0$ . The slepton masses are determined from the GUT relations without stau mixing. See their Fig. 1 for the dependence of the limit on  $\Delta m$ .
- <sup>2</sup> ABBIENDI 06b use 600 pb<sup>-1</sup> of data from  $\sqrt{s}=189\text{--}209$  GeV. They look for events from pair-produced staus in a GMSB scenario with  $\tilde{t}$  co-NLSP including prompt  $\tilde{t}$  decays to dileptons +  $\cancel{E}$  final states, large impact parameters, kinked tracks and heavy stable charged particles. Limits on the cross-section are computed as a function of  $m(\tilde{t})$  and the lifetime, see their Fig. 7. The limit is compared to the  $\sigma \cdot BR^2$  from a scan over the GMSB parameter space. The highest mass limit is reached for  $\tilde{\mu}_R$ , from which the quoted mass limit is derived by subtracting  $m_{\tilde{\tau}}$ .
- <sup>3</sup> ABDALLAH 03D use data from  $\sqrt{s}=130\text{--}208$  GeV to search for tracks with large impact parameter or visible decay vertices and for heavy charged stable particles. Limits are obtained as function of  $m(\tilde{G})$ , after combining these results with the search for slepton pair production in the SUGRA framework from ABDALLAH 03M to cover prompt decays. The above limit is reached for prompt decays and assumes the degeneracy of the sleptons. For limits at different  $m(\tilde{G})$ , see their Fig. 9. Supersedes the results of ABREU 01G.
- <sup>4</sup> ACHARD 02 searches for the production of sparticles in the case of  $\tilde{R}$  prompt decays with  $L\tilde{L}\tilde{E}$  or  $U\tilde{D}\tilde{D}$  couplings at  $\sqrt{s}=189\text{--}208$  GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale and no mixing in the slepton sector, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit holds for  $L\tilde{L}\tilde{E}$  couplings and increases to 88.7 GeV for  $U\tilde{D}\tilde{D}$  couplings. For L3 limits from  $L\tilde{Q}\tilde{D}$  couplings, see ACCIARRI 01.
- <sup>5</sup> ABBIENDI 01 looked for final states with  $\gamma\gamma\tilde{E}$ ,  $\ell\ell\tilde{E}$ , with possibly additional activity and four leptons +  $\cancel{E}$  to search for prompt decays of  $\tilde{\chi}_1^0$  or  $\tilde{\ell}_1$  in GMSB. They derive limits in the plane  $(m_{\tilde{\chi}_1^0}, m_{\tilde{\ell}_1})$ , see Fig. 6, allowing either the  $\tilde{\chi}_1^0$  or a  $\tilde{\ell}_1$  to be the NLSP. Two scenarios are considered:  $\tan\beta=2$  with the 3 sleptons degenerate in mass and  $\tan\beta=20$  where the  $\tilde{\tau}_1$  is lighter than the other sleptons. Data taken at  $\sqrt{s}=189$  GeV. For  $\tan\beta=20$ , the obtained limits are  $m_{\tilde{\tau}_1} > 69$  GeV and  $m_{\tilde{e}_1, \tilde{\mu}_1} > 88$  GeV.
- <sup>6</sup> ABREU 01 looked for acoplanar dilepton + diphoton +  $\cancel{E}$  final states from  $\tilde{e}$  cascade decays at  $\sqrt{s}=130\text{--}189$  GeV. See Fig. 9 for limits on the  $(\mu, M_2)$  plane for  $m_{\tilde{e}}=80$  GeV,  $\tan\beta=1.0$ , and assuming degeneracy of  $\tilde{\mu}$  and  $\tilde{e}$ .
- <sup>7</sup> ACCIARRI 01 searches for multi-lepton and/or multi-jet final states from  $\tilde{R}$  prompt decays with  $L\tilde{L}\tilde{E}$ ,  $L\tilde{Q}\tilde{D}$ , or  $U\tilde{D}\tilde{D}$  couplings at  $\sqrt{s}=189$  GeV. The search is performed for direct and indirect decays of neutralinos, charginos, and scalar leptons, with the  $\tilde{\chi}_1^0$  or a  $\tilde{\ell}$  as LSP and assuming one coupling to be nonzero at a time. Mass limits are derived using simultaneously the constraints from the neutralino, chargino, and slepton analyses; and the  $Z^0$  width measurements from ACCIARRI 00c in a scan of the parameter space assuming MSUGRA with gaugino and scalar mass universality. Updates and supersedes the results from ACCIARRI 99i.
- <sup>8</sup> ABREU 00v use data from  $\sqrt{s}=130\text{--}189$  GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of  $m_{\tilde{\tau}}$ , after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00q. For limits at different  $m_{\tilde{G}}$ , see their Fig. 12. The above limit assumes the degeneracy of stau and smuon.

### Long-lived $\tilde{\tau}$ (Slepton) MASS LIMIT

Limits on scalar leptons which leave detector before decaying. Limits from Z decays are independent of lepton flavor. Limits from continuum  $e^+e^-$  annihilation are also independent of flavor for smuons and staus. Selection limits from  $e^+e^-$  collisions in the continuum depend on MSSM parameters because of the additional neutralino exchange contribution.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>440	95	<sup>1</sup> AAD	15AE ATLS	mGMSB, $M_{mess}=250$ TeV, $N_5=3$ , $\mu > 0$ , $C_{grav}=5000$ , $\tan\beta=10$
>385	95	<sup>1</sup> AAD	15AE ATLS	mGMSB, $M_{mess}=250$ TeV, $N_5=3$ , $\mu > 0$ , $C_{grav}=5000$ , $\tan\beta=50$
>286	95	<sup>1</sup> AAD	15AE ATLS	direct $\tilde{\tau}$ production
none 124–309	95	<sup>2</sup> AAIJ	15BD LHCb	long-lived $\tilde{\tau}$ , mGMSB, SPS7
> 98	95	<sup>3</sup> ABBIENDI	03L OPAL	$\tilde{\mu}_R, \tilde{\tau}_R$
none 2–87.5	95	<sup>4</sup> ABREU	00q DLPH	$\tilde{\mu}_R, \tilde{\tau}_R$
> 81.2	95	<sup>5</sup> ACCIARRI	99H L3	$\tilde{\mu}_R, \tilde{\tau}_R$
> 81	95	<sup>6</sup> BARATE	98K ALEP	$\tilde{\mu}_R, \tilde{\tau}_R$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>300	95	<sup>7</sup> AAD	13AA ATLS	long-lived $\tilde{\tau}$ , GMSB, $\tan\beta=5\text{--}20$
		<sup>8</sup> ABZOV	13B D0	long-lived $\tilde{\tau}$ , $100 < m_{\tilde{\tau}} < 300$ GeV
>339	95	<sup>9,10</sup> CHATRCHYAN	13AB CMS	long-lived $\tilde{\tau}$ , direct $\tilde{\tau}_1$ pair prod., minimal GMSB, SPS line 7
>500	95	<sup>9,11</sup> CHATRCHYAN	13AB CMS	long-lived $\tilde{\tau}$ , $\tilde{\tau}_1$ from direct pair prod. and from decay of heavier SUSY particles, minimal GMSB, SPS line 7
>314	95	<sup>12</sup> CHATRCHYAN	12L CMS	long-lived $\tilde{\tau}$ , $\tilde{\tau}_1$ from decay of heavier SUSY particles, minimal GMSB, SPS line 7
>136	95	<sup>13</sup> AAD	11P ATLS	stable $\tilde{\tau}$ , GMSB scenario, $\tan\beta=5$

- <sup>1</sup> AAD 15AE searched in 19.1 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=8$  TeV for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ALTA muon system. In the absence of an excess of events above the expected backgrounds, limits are set on stable  $\tilde{\tau}$  sleptons in various scenarios, see Figs. 5–7.
- <sup>2</sup> AAIJ 15BD searched in 3.0 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=7$  and 8 TeV for evidence of Drell-Yan pair production of long-lived  $\tilde{\tau}$  particles. No evidence for such particles is observed and 95% C.L. upper limits on the cross section of  $\tilde{\tau}$  pair production are derived, see Fig. 7. In the mGMSB, assuming the SPS7 benchmark scenario  $\tilde{\tau}$  masses between 124 and 309 GeV are excluded at 95% C.L.

- <sup>3</sup> ABBIENDI 03L used  $e^+e^-$  data at  $\sqrt{s}=130\text{--}209$  GeV to select events with two high momentum tracks with anomalous  $dE/dx$ . The excluded cross section is compared to the theoretical expectation as a function of the heavy particle mass in their Fig. 3. The limit improves to 98.5 GeV for  $\tilde{\mu}_L$  and  $\tilde{\tau}_L$ . The bounds are valid for colorless spin 0 particles with lifetimes longer than  $10^{-6}$  s. Supersedes the results from ACKERSTAFF 98P.
- <sup>4</sup> ABREU 00q searches for the production of pairs of heavy, charged stable particles in  $e^+e^-$  annihilation at  $\sqrt{s}=130\text{--}189$  GeV. The upper bound improves to 88 GeV for  $\tilde{\mu}_L, \tilde{\tau}_L$ . These limits include and update the results of ABREU 98P.
- <sup>5</sup> ACCIARRI 99H searched for production of pairs of back-to-back heavy charged particles at  $\sqrt{s}=130\text{--}183$  GeV. The upper bound improves to 82.2 GeV for  $\tilde{\mu}_L, \tilde{\tau}_L$ .
- <sup>6</sup> The BARATE 98K mass limit improves to 82 GeV for  $\tilde{\mu}_L, \tilde{\tau}_L$ . Data collected at  $\sqrt{s}=161\text{--}184$  GeV.
- <sup>7</sup> AAD 13AA searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=7$  TeV for events containing long-lived massive particles in a GMSB framework. No significant excess above the expected background was found. A 95% C.L. lower limit of 300 GeV is placed on long-lived  $\tilde{\tau}$ 's in the GMSB model with  $M_{mess}=250$  TeV,  $N_5=3$ ,  $\mu > 0$ , for  $\tan\beta=5\text{--}20$ . The lower limit on the GMSB breaking scale  $\Lambda$  was found to be 99–110 TeV, for  $\tan\beta$  values between 5 and 40, see Fig. 4 (top). Also, directly produced long-lived sleptons, or sleptons decaying to long-lived ones, are excluded at 95% C.L. up to a  $\tilde{\tau}$  mass of 278 GeV for models with slepton splittings smaller than 50 GeV.
- <sup>8</sup> ABZOV 13b looked in 6.3 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV for charged massive long-lived particles in events with muon-like particles that have both speed and ionization energy loss inconsistent with muons produced in beam collisions. In the absence of an excess, limits are set at 95% C.L. on the production cross section of stau leptons in the mass range 100–300 GeV, see their Table 20 and Fig. 23.
- <sup>9</sup> CHATRCHYAN 13AB looked in 5.0 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=7$  TeV and in 18.8 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=8$  TeV for events with heavy stable particles, identified by their anomalous  $dE/dx$  in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of  $\tilde{\tau}_1$ 's. No evidence for an excess over the expected background is observed. Supersedes CHATRCHYAN 12L.
- <sup>10</sup> CHATRCHYAN 13AB limits are derived for pair production of  $\tilde{\tau}_1$  as a function of mass in minimal GMSB scenarios along the Snowmass Points and Slopes (SPS) line 7 (see Fig. 8 and Table 7). The limit given here is valid for direct pair  $\tilde{\tau}_1$  production.
- <sup>11</sup> CHATRCHYAN 13AB limits are derived for the production of  $\tilde{\tau}_1$  as a function of mass in minimal GMSB scenarios along the Snowmass Points and Slopes (SPS) line 7 (see Fig. 8 and Table 7). The limit given here is valid for the production of  $\tilde{\tau}_1$  from both direct pair production and from the decay of heavier supersymmetric particles.
- <sup>12</sup> CHATRCHYAN 12L looked in 5.0 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=7$  TeV for events with heavy stable particles, identified by their anomalous  $dE/dx$  in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of  $\tilde{\tau}_1$ 's. No evidence for an excess over the expected background is observed. Limits are derived for the production of  $\tilde{\tau}_1$  as a function of mass in minimal GMSB scenarios along the Snowmass Points and Slopes (SPS) line 7 (see Fig. 3). The limit given here is valid for the production of  $\tilde{\tau}_1$  in the decay of heavier supersymmetric particles.
- <sup>13</sup> AAD 11P looked in 37 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s}=8$  TeV for events with two heavy stable particles, reconstructed in the Inner tracker and the Muon System and identified by their time of flight in the Muon System. No evidence for an excess over the SM expectation is observed. Limits on the mass are derived, see Fig. 3, for  $\tilde{\tau}$  in a GMSB scenario and for sleptons produced by electroweak processes only, in which case the limit degrades to 110 GeV.

### $\tilde{q}$ (Squark) MASS LIMIT

For  $m_{\tilde{q}} > 60\text{--}70$  GeV, it is expected that squarks would undergo a cascade decay via a number of neutralinos and/or charginos rather than undergo a direct decay to photinos as assumed by some papers. Limits obtained when direct decay is assumed are usually higher than limits when cascade decays are included.

Limits from  $e^+e^-$  collisions depend on the mixing angle of the lightest mass eigenstate  $\tilde{q}_1 = \tilde{q}_R \sin\theta_q + \tilde{q}_L \cos\theta_q$ . It is usually assumed that only the sbottom and stop squarks have non-trivial mixing angles (see the stop and sbottom sections). Here, unless otherwise noted, squarks are always taken to be either left/right degenerate, or purely of left or right type. Data from Z decays have set squark mass limits above 40 GeV, in the case of  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  decays if  $\Delta m = m_{\tilde{q}} - m_{\tilde{\chi}_1^0} \gtrsim 5$  GeV. For smaller values of  $\Delta m$ , current constraints on the invisible width of the Z ( $\Delta\Gamma_{inv} < 2.0$  MeV, LEP 00) exclude  $m_{\tilde{\mu}_L, R} < 44$  GeV,  $m_{\tilde{d}_R} < 33$  GeV,  $m_{\tilde{d}_L} < 44$  GeV and, assuming all squarks degenerate,  $m_{\tilde{q}} < 45$  GeV.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;1450 (CL = 95%) OUR EVALUATION</b>				
> 850	95	<sup>1</sup> AAD	15BV ATLS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 100$ GeV
> 250	95	<sup>2</sup> AAD	15CS ATLS	photon + $\cancel{E}_T$ , $pp \rightarrow \tilde{q}\tilde{q}^*\gamma$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , $m_{\tilde{q}} - m_{\tilde{\chi}_1^0} = m_c$
> 490	95	<sup>3</sup> AAD	15K ATLS	$\tilde{c} \rightarrow c\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} < 200$ GeV
> 875	95	<sup>4</sup> KHACHATRY...	15AF CMS	$\tilde{g} \rightarrow q\tilde{\chi}_1^0$ , simplified model, 8 degenerate light $\tilde{q}$ , $m_{\tilde{\chi}_1^0} = 0$
> 520	95	<sup>4</sup> KHACHATRY...	15AF CMS	$\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , simplified model, single light squark, $m_{\tilde{\chi}_1^0} = 0$
>1450	95	<sup>4</sup> KHACHATRY...	15AF CMS	CMSSM, $\tan\beta=30$ , $A_0=0$ , $-2\max(m_0, m_{1/2})$ , $\mu > 0$
> 850	95	<sup>5</sup> AAD	14AE ATLS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ simplified model, mass degenerate first and second generation squarks, $m_{\tilde{\chi}_1^0} = 0$ GeV

See key on page 601

# Searches Particle Listings

## Supersymmetric Particle Searches

> 440	95	<sup>5</sup> AAD	14AE ATLS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ simplified model, single light-flavour squark, $m_{\tilde{\chi}_1^0} = 0$ GeV	1
>1700	95	<sup>5</sup> AAD	14AE ATLS	jets + $\cancel{E}_T$ , mSUGRA/CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$	2
> 800	95	<sup>6</sup> CHATRCHYAN14AH	CMS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50$ GeV	3
> 780	95	<sup>7</sup> CHATRCHYAN14I	CMS	multijets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 200$ GeV	4
>1360	95	<sup>8</sup> AAD	13L ATLS	jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{g}} = m_{\tilde{q}}$	5
>1200	95	<sup>9</sup> AAD	13Q ATLS	$\gamma + b + \cancel{E}_T$ , higgsino-like neutralino, $m_{\tilde{\chi}_1^0} > 220$ GeV, GMSB	6
>1250	95	<sup>10</sup> CHATRCHYAN13	CMS	$\ell^\pm \ell^\mp$ + jets + $\cancel{E}_T$ , CMSSM	7
>1430	95	<sup>11</sup> CHATRCHYAN13G	CMS	0,1,2, $\geq 3$ b-jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$	8
> 750	95	<sup>12</sup> CHATRCHYAN13H	CMS	$2\gamma + \geq 4$ jets + low $\cancel{E}_T$ , stealth SUSY model	9
> 820	95	<sup>13</sup> CHATRCHYAN13T	CMS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 0$ GeV	10
>1200	95	<sup>14</sup> AAD	12AX ATLS	$\ell +$ jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$	11
> 870	95	<sup>15</sup> AAD	12CJ ATLS	$\ell^\pm$ + jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$	12
> 950	95	<sup>16</sup> AAD	12CP ATLS	$2\gamma + \cancel{E}_T$ , GMSB, bino NLSP, $m_{\tilde{\chi}_1^0} > 50$ GeV	13
> 760	95	<sup>17</sup> AAD	12W ATLS	jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$	14
>1110	95	<sup>18</sup> CHATRCHYAN12	CMS	e, $\mu$ , jets, razor, CMSSM	15
>1180	95	<sup>19</sup> CHATRCHYAN12AE	CMS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} < 200$ GeV	16
• • • We do not use the following data for averages, fits, limits, etc. • • •		<sup>20</sup> CHATRCHYAN12AL	CMS	$\geq 3\ell^\pm$ , $R$	17
>1650	95	<sup>21</sup> AAD	15AI ATLS	jets + $\cancel{E}_T$ , CMSSM	18
> 790	95	<sup>22</sup> AAD	15BV ATLS	jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$	19
> 820	95	<sup>1</sup> AAD	15BV ATLS	jets + $\cancel{E}_T$ , $\tilde{q} \rightarrow qW\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 100$ GeV	20
> 850	95	<sup>1</sup> AAD	15BV ATLS	2 or 3 leptons + jets, $\tilde{q}$ decays via sleptons, $m_{\tilde{\chi}_1^0} = 100$ GeV	21
>1000	95	<sup>1</sup> AAD	15BV ATLS	$\tau$ , $\tilde{q}$ decays via staus, $m_{\tilde{\chi}_1^0} = 50$ GeV	22
> 700	95	<sup>23</sup> AAD	15CB ATLS	jets, $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , $\tilde{\chi}_1^0 \rightarrow \ell q$ , RPV, $m_{\tilde{\chi}_1^0} = 108$ GeV and $2.5 < c\tau_{\tilde{\chi}_1^0} < 200$ mm	23
> 550	95	<sup>24</sup> KHACHATRY...15AR	CMS	$\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , $\tilde{\chi}_1^0 \rightarrow \tilde{S}\tilde{g}$ , $\tilde{S} \rightarrow \tilde{S}\tilde{G}$ , $\tilde{S} \rightarrow g\tilde{g}$ , $m_{\tilde{S}} = 100$ GeV, $m_{\tilde{S}} = 90$ GeV	24
>1500	95	<sup>25</sup> KHACHATRY...15AZ	CMS	$\ell^\pm$ , $\tilde{q} \rightarrow q\tilde{\chi}_1^\pm$ , $\tilde{\chi}_1^\pm \rightarrow \tilde{S}W^\pm$ , $\tilde{S} \rightarrow \tilde{S}\tilde{G}$ , $\tilde{S} \rightarrow g\tilde{g}$ , $m_{\tilde{S}} = 100$ GeV, $m_{\tilde{S}} = 90$ GeV	25
>1000	95	<sup>25</sup> KHACHATRY...15AZ	CMS	$\geq 2\gamma$ , $\geq 1$ jet, (Razor), bino-like NLSP, $m_{\tilde{\chi}_1^0} = 375$ GeV	26
> 670	95	<sup>26</sup> AAD	14E ATLS	$\ell^\pm \ell^\mp$ ( $\ell^\mp$ ) + jets, $\tilde{q} \rightarrow q'\tilde{\chi}_1^\pm$ , $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\tilde{\chi}_2^0$ , $\tilde{\chi}_2^0 \rightarrow Z^{(*)}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 300$ GeV	27
> 780	95	<sup>26</sup> AAD	14E ATLS	$\ell^\pm \ell^\mp$ ( $\ell^\mp$ ) + jets, $\tilde{q} \rightarrow q'\tilde{\chi}_1^\pm/\tilde{\chi}_2^0$ , $\tilde{\chi}_1^\pm \rightarrow \ell^\pm \nu \tilde{\chi}_1^0$ , $\tilde{\chi}_2^0 \rightarrow \ell^\pm \ell^\mp (\nu\nu)\tilde{\chi}_1^0$ simplified model	28
> 700	95	<sup>27</sup> CHATRCHYAN13AO	CMS	$\ell^\pm \ell^\mp$ + jets + $\cancel{E}_T$ , CMSSM, $m_0 < 700$ GeV	29
>1350	95	<sup>28</sup> CHATRCHYAN13AV	CMS	jets (+ leptons) + $\cancel{E}_T$ , CMSSM, $m_{\tilde{g}} = m_{\tilde{q}}$	30
> 800	95	<sup>29</sup> CHATRCHYAN13W	CMS	$\geq 1$ photons + jets + $\cancel{E}_T$ , GGM, wino-like NLSP, $m_{\tilde{\chi}_1^0} = 375$ GeV	31
>1000	95	<sup>29</sup> CHATRCHYAN13W	CMS	$\geq 2$ photons + jets + $\cancel{E}_T$ , GGM, bino-like NLSP, $m_{\tilde{\chi}_1^0} = 375$ GeV	32
> 340	95	<sup>30</sup> DREINER	12A THEO	$m_{\tilde{q}} \sim m_{\tilde{\chi}_1^0}$	33
> 650	95	<sup>31</sup> DREINER	12A THEO	$m_{\tilde{q}} = m_{\tilde{g}} \sim m_{\tilde{\chi}_1^0}$	34

- <sup>1</sup> AAD 15BV summarized and extended ATLAS searches for gluinos and first- and second-generation squarks in final states containing jets and missing transverse momentum, with or without leptons or  $b$ -jets in the  $\sqrt{s} = 8$  TeV data set collected in 2012. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on the squark mass in several  $R$ -parity conserving models. See their Figs. 9, 11, 18, 22, 24, 27, 28.
- <sup>2</sup> AAD 15CS searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for evidence of pair production of squarks, decaying into a quark and a neutralino, where a photon was radiated either from an initial-state quark, from an intermediate squark, or from a final-state quark. No evidence was found for an excess above the expected level of Standard Model background and a 95% C.L. exclusion limit was set on the squark mass as a function of the squark-neutralino mass difference, see Fig. 19.
- <sup>3</sup> AAD 15K searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing at least two jets, where the two leading jets are each identified as originating from  $c$ -quarks, and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the mass of superpartners of charm quarks ( $\tilde{c}$ ). Assuming that the decay  $\tilde{c} \rightarrow c\tilde{\chi}_1^0$  takes place 100% of the time, a scalar charm mass below 490 GeV is excluded for  $m_{\tilde{\chi}_1^0} < 200$  GeV. For more details, see their Fig. 2.
- <sup>4</sup> KHACHATRYAN 15AF searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least two energetic jets and significant  $\cancel{E}_T$ , using the transverse mass variable  $M_{T2}$  to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the squark mass in simplified models where the decay  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, both for the case of a single light squark or 8 degenerate squarks, see Fig. 12. See also Table 5. Exclusions in the CMSSM, assuming  $\tan\beta = 30$ ,  $A_0 = -2 \max(m_0, m_{1/2})$  and  $\mu > 0$ , are also presented, see Fig. 15.
- <sup>5</sup> AAD 14AE searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for strongly produced supersymmetric particles in events containing jets and large missing transverse momentum, and no electrons or muons. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing squarks that decay via  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , where either a single light state or two degenerate generations of squarks are assumed, see Fig. 10.
- <sup>6</sup> CHATRCHYAN 14AH searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with at least two energetic jets and significant  $\cancel{E}_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 28. Exclusions in the CMSSM, assuming  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , are also presented, see Fig. 26.
- <sup>7</sup> CHATRCHYAN 14I searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing multijets and large  $\cancel{E}_T$ . No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing squarks that decay via  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$ , where either a single light state or two degenerate generations of squarks are assumed, see Fig. 7a.
- <sup>8</sup> AAD 13L searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for the production of squarks and gluinos in events containing jets, missing transverse momentum and no high- $p_T$  electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , squarks and gluinos of equal mass are excluded for masses below 1360 GeV at 95% C.L. In a simplified model containing only squarks of the first two generations, a gluino octet and a massless neutralino, squark masses below 1320 GeV are excluded at 95% C.L. for gluino masses below 2 TeV. See Figures 10–15 for more precise bounds.
- <sup>9</sup> AAD 13Q searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events containing a high- $p_T$  isolated photon, at least one jet identified as originating from a bottom quark, and high missing transverse momentum. Such signatures may originate from supersymmetric models with gauge-mediated supersymmetry breaking in events in which one of a pair of higgsino-like neutralinos decays into a photon and a gravitino while the other decays into a Higgs boson and a gravitino. No significant excess above the expected background was found and limits were set on the squark mass as a function of the neutralino mass in a generalized GMSB model (GGM) with a higgsino-like neutralino NLSP, see their Fig. 4. For neutralino masses greater than 220 GeV, squark masses below 1020 GeV are excluded at 95% C.L.
- <sup>10</sup> CHATRCHYAN 13 looked in 4.98 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with two opposite-sign leptons ( $e, \mu, \tau$ ), jets and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the mSUGRA/CMSSM model with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , see Fig. 6.
- <sup>11</sup> CHATRCHYAN 13G searched in 4.98 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for the production of squarks and gluinos in events containing 0,1,2,  $\geq 3$   $b$ -jets, missing transverse momentum and no electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with  $\tan\beta = 10$ ,  $A_0 = 0$ , and  $\mu > 0$ , squarks and gluinos of equal mass are excluded for masses below 1250 GeV at 95% C.L. Exclusions are also derived in various simplified models, see Fig. 7.
- <sup>12</sup> CHATRCHYAN 13H searched in 4.96 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with two photons,  $\geq 4$  jets and low  $\cancel{E}_T$  due to  $\tilde{q} \rightarrow \gamma\tilde{\chi}_1^0$  decays in a stealth SUSY framework, where the  $\tilde{\chi}_1^0$  decays through a singlino ( $\tilde{S}$ ) intermediate state to  $\gamma S \tilde{G}$ , with the singlet state  $S$  decaying to two jets. No significant excess above the expected background was found and limits were set in a particular  $R$ -parity conserving stealth SUSY model. The model assumes  $m_{\tilde{\chi}_1^0} = 0.5 m_{\tilde{q}}$ ,  $m_{\tilde{S}} = 100$  GeV and  $m_S = 90$  GeV. Under these assumptions, squark masses less than 1430 GeV were excluded at the 95% C.L.
- <sup>13</sup> CHATRCHYAN 13T searched in 11.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least two energetic jets and significant  $\cancel{E}_T$ , using the  $\alpha_T$  variable to discriminate between processes with genuine and misreconstructed  $\cancel{E}_T$ . No significant excess above the Standard Model expectations is observed. Limits are set on squark masses in simplified models where the decay  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, assuming an eightfold degeneracy of the masses of the first two generation squarks, see Fig. 8 and Table 9. Also limits in the case of a single light squark are given.
- <sup>14</sup> AAD 12AX searched in 1.04 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for supersymmetry in events containing jets, missing transverse momentum and one isolated electron or muon. No excess over the expected SM background is observed and model-independent limits are set on the cross section of new physics contributions to the signal regions. In mSUGRA/CMSSM models with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , squarks and gluinos of

# Searches Particle Listings

## Supersymmetric Particle Searches

- equal mass are excluded for masses below 820 GeV at 95% C.L. Limits are also set on simplified models for squark production and decay via an intermediate chargino and on supersymmetric models with bilinear  $R$ -parity violation. Supersedes AAD 11c.
- 15 AAD 12CJ searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events containing one or more isolated leptons (electrons or muons), jets and  $E_{T\text{miss}}$ . The observations are in good agreement with the SM expectations and exclusion limits have been set in number of SUSY models. In the mSUGRA/CMSSM model with  $\tan\beta = 10$ ,  $A_0 = 0$ , and  $\mu > 0$ , 95% C.L. exclusion limits have been derived for  $m_{\tilde{q}} < 1200 \text{ GeV}$ , assuming equal squark and gluino masses. In minimal GMSB, values of the effective SUSY breaking scale  $\Lambda < 50 \text{ TeV}$  are excluded at 95% C.L. for  $\tan\beta < 45$ . Also exclusion limits in a number of simplified models have been presented, see Figs. 10 and 12.
- 16 AAD 12CP searched in  $4.8 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with two photons and large  $E_{T\text{miss}}$  due to  $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$  decays in a GMSB framework. No significant excess above the expected background was found and limits were set on the squark mass as a function of the neutralino mass in a generalized GMSB model (GGM) with a bino-like neutralino NLSP. The other sparticle masses were decoupled,  $\tan\beta = 2$  and  $\tau_{\text{NLSP}} < 0.1 \text{ mm}$ . Also, in the framework of the SP58 model, a 95% C.L. lower limit was set on the breaking scale  $\Lambda$  of 196 TeV.
- 17 AAD 12W searched in  $1.04 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for the production of squarks and gluinos in events containing jets, missing transverse momentum and no electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , squarks and gluinos of equal mass are excluded for masses below 950 GeV at 95% C.L. In a simplified model containing only squarks of the first two generations, a gluino octet and a massless neutralino, squark masses below 875 GeV are excluded at 95% C.L.
- 18 CHATRCHYAN 12 looked in  $35 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with  $e$  and/or  $\mu$  and/or jets, a large total transverse energy, and  $E_{T\text{miss}}$ . The event selection is based on the dimensionless razor variable  $R$ , related to the  $E_{T\text{miss}}$  and  $M_{\text{RP}}$ , an indicator of the heavy particle mass scale. No evidence for an excess over the expected background is observed. Limits are derived in the CMSSM ( $m_0, m_{1/2}$ ) plane for  $\tan\beta = 3, 10$  and 50 (see Fig. 7 and 8). Limits are also obtained for Simplified Model Spectra.
- 19 CHATRCHYAN 12AE searched in  $4.98 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with at least three jets and large missing transverse momentum. No significant excesses over the expected SM backgrounds are observed and 95% C.L. limits on the production cross section of squarks in a scenario where  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  with a 100% branching ratio, see Fig. 3. For  $m_{\tilde{\chi}_1^0} < 200 \text{ GeV}$ , values of  $m_{\tilde{q}}$  below 760 GeV are excluded at 95% C.L. Also limits in the CMSSM are presented, see Fig. 2.
- 20 CHATRCHYAN 12AL looked in  $4.98 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for anomalous production of events with three or more isolated leptons. Limits on squark and gluino masses are set in  $R$  SUSY models with leptonic  $LL\tilde{E}$  couplings,  $\lambda_{123} > 0.05$ , and hadronic  $U\tilde{D}\tilde{D}$  couplings,  $\lambda_{112}'' > 0.05$ , see their Fig. 5. In the  $U\tilde{D}\tilde{D}$  case the leptons arise from supersymmetric cascade decays. A very specific supersymmetric spectrum is assumed. All decays are prompt.
- 21 CHATRCHYAN 12AT searched in  $4.73 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for the production of squarks and gluinos in events containing jets, missing transverse momentum and no electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , squarks with masses below 1110 GeV are excluded at 95% C.L. Squarks and gluinos of equal mass are excluded for masses below 1180 GeV at 95% C.L. Exclusions are also derived in various simplified models, see Fig. 6.
- 22 AAD 15AI searched in  $20 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events containing at least one isolated lepton (electron or muon), jets, and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the squark masses in the CMSSM/mSUGRA, see Fig. 15, in the NUHMG, see Fig. 16, and in various simplified models, see Figs. 19–21.
- 23 AAD 15CB searched for events containing at least one long-lived particle that decays at a significant distance from its production point (displaced vertex, DV) into two leptons or into five or more charged particles in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$ . The dilepton signature is characterised by DV formed from at least two lepton candidates. Four different final states were considered for the multitrack signature, in which the DV must be accompanied by a high-transverse momentum muon or electron candidate that originates from the DV, jets or missing transverse momentum. No events were observed in any of the signal regions. Results were interpreted in SUSY scenarios involving  $R$ -parity violation, split supersymmetry, and gauge mediation. See their Fig. 14–20.
- 24 KHACHATRYAN 15AR searched in  $19.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events containing jets, either a charged lepton or a photon, and low missing transverse momentum. No significant excess above the Standard Model expectations is observed. Limits are set on the squark mass in a stealth SUSY model where the decays  $\tilde{q} \rightarrow q\tilde{\chi}_1^\pm$ ,  $\tilde{\chi}_1^\pm \rightarrow \tilde{S}W^\pm$ ,  $\tilde{S} \rightarrow S\tilde{G}$  and  $S \rightarrow g\tilde{g}$ , with  $m_{\tilde{S}} = 100 \text{ GeV}$  and  $m_{\tilde{S}} = 90 \text{ GeV}$ , take place with a branching ratio of 100%. See Fig. 6 for  $\gamma$  or Fig. 7 for  $e^\pm$  analyses.
- 25 KHACHATRYAN 15AZ searched in  $19.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with either at least one photon, hadronic jets and  $E_{T\text{miss}}$  (single photon channel) or with at least two photons and at least one jet and using the razor variables. No significant excess above the Standard Model expectations is observed. Limits are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for both a bino-like and wino-like neutralino NLSP scenario, see Fig. 8 and 9.
- 26 AAD 14E searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for strongly produced supersymmetric particles in events containing jets and two same-sign leptons or three leptons. The search also utilises jets originating from  $b$ -quarks, missing transverse momentum and other variables. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos and squarks, see Figs. 5 and 6. In the  $\tilde{q} \rightarrow q'\tilde{\chi}_1^\pm$ ,  $\tilde{\chi}_1^\pm \rightarrow W(*)\tilde{\chi}_2^0$ ,  $\tilde{\chi}_2^0 \rightarrow Z(*)\tilde{\chi}_1^0$  simplified model, the following assumptions have been made:  $m_{\tilde{\chi}_1^\pm} = 0.5 m_{\tilde{\chi}_1^0} + m_{\tilde{g}}$ ,  $m_{\tilde{\chi}_2^0} = 0.5 (m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_1^\pm})$ . In the  $\tilde{q} \rightarrow q'\tilde{\chi}_1^\pm$  or  $\tilde{q} \rightarrow q'\tilde{\chi}_2^0$ ,  $\tilde{\chi}_1^\pm \rightarrow e^\pm\nu\tilde{\chi}_1^0$  or  $\tilde{\chi}_2^0 \rightarrow e^\pm e^\mp(\nu\nu)\tilde{\chi}_1^0$  simplified model, the following assumptions have been made:  $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0} = 0.5 (m_{\tilde{\chi}_1^0} + m_{\tilde{q}})$ ,  $m_{\tilde{\chi}_1^0} < 460 \text{ GeV}$ . Limits are also derived in the mSUGRA/CMSSM, BRPV and GMSB models, see their Fig. 8.
- 27 CHATRCHYAN 2013AO searched in  $4.98 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with two opposite-sign isolated leptons accompanied by hadronic jets and  $E_{T\text{miss}}$ . No significant excesses over the expected SM backgrounds are observed and 95% C.L.

exclusion limits are derived in the mSUGRA/CMSSM model with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , see Fig. 8.

- 28 CHATRCHYAN 13AV searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for new heavy particle pairs decaying into jets (possibly  $b$ -tagged), leptons and  $E_{T\text{miss}}$  using the Razor variables. No significant excesses over the expected SM backgrounds are observed and 95% C.L. exclusion limits are derived in the mSUGRA/CMSSM model with  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , see Fig. 3. The results are also interpreted in various simplified models, see Fig. 4.
- 29 CHATRCHYAN 13W searched in  $4.93 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with one or more photons, hadronic jets and  $E_{T\text{miss}}$ . No significant excess above the Standard Model expectations is observed. Limits are set on squark masses in the general gauge-mediated SUSY breaking model (GGM), for both a wino-like and bino-like neutralino NLSP scenario, see Fig. 5.
- 30 DREINER 12A reassesses constraints from CMS (at  $7 \text{ TeV}$ ,  $\sim 4.4 \text{ fb}^{-1}$ ) under the assumption that the first and second generation squarks and the lightest SUSY particle are quasi-degenerate in mass (compressed spectrum).
- 31 DREINER 12A reassesses constraints from CMS (at  $7 \text{ TeV}$ ,  $\sim 4.4 \text{ fb}^{-1}$ ) under the assumption that the first and second generation squarks, the gluino, and the lightest SUSY particle are quasi-degenerate in mass (compressed spectrum).

### Long-lived $\tilde{q}$ (Squark) MASS LIMIT

The following are bounds on long-lived scalar quarks, assumed to hadronise into hadrons with lifetime long enough to escape the detector prior to a possible decay. Limits may depend on the mixing angle of mass eigenstates:  $\tilde{q}_1 = \tilde{q}_L \cos\theta_q + \tilde{q}_R \sin\theta_q$ .

The coupling to the  $Z^0$  boson vanishes for up-type squarks when  $\theta_q = 0.98$ , and for down type squarks when  $\theta_q = 1.17$ .

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 845	95	1 AAD	15AE ATLS	$\tilde{b}$ R-hadron, stable, Regge model
> 900	95	1 AAD	15AE ATLS	$\tilde{t}$ R-hadron, stable, Regge model
>1500	95	1 AAD	15AE ATLS	$\tilde{g}$ decaying to 300 GeV stable sleptons, LeptoSUSY model
> 751	95	2 AAD	15BM ATLS	$\tilde{b}$ R-hadron, stable, Regge model
> 766	95	2 AAD	15BM ATLS	$\tilde{t}$ R-hadron, stable, Regge model
> 525	95	3 KHACHATRYAN	15AK CMS	$\tilde{g}$ R-hadrons, $10 \mu\text{s} < \tau < 1000 \text{ s}$
> 470	95	3 KHACHATRYAN	15AK CMS	$\tilde{g}$ R-hadrons, $1 \mu\text{s} < \tau < 1000 \text{ s}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> 683	95	4 AAD	13AA ATLS	$\tilde{t}$ , R-hadrons, generic interaction model
> 612	95	5 AAD	13AA ATLS	$\tilde{b}$ , R-hadrons, generic interaction model
> 344	95	6 AAD	13BC ATLS	R-hadrons, $\tilde{t} \rightarrow b\tilde{\chi}_1^0$ , Regge model, lifetime between $10^{-5}$ and $10^3 \text{ s}$ , $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
> 379	95	7 AAD	13BC ATLS	R-hadrons, $\tilde{t} \rightarrow t\tilde{\chi}_1^0$ , Regge model, lifetime between $10^{-5}$ and $10^3 \text{ s}$ , $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
> 935	95	8 CHATRCHYAN	13AB CMS	long-lived $\tilde{t}$ forming R-hadrons, cloud interaction model

- 1 AAD 15AE searched in  $19.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ATLAS muon system. In the absence of an excess of events above the expected backgrounds, limits are set on R-hadrons in various scenarios, see Fig. 11. Limits are also set in LeptoSUSY models where the gluino decays to stable 300 GeV leptons, see Fig. 9.
- 2 AAD 15BM searched in  $18.4 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for stable and metastable non-relativistic charged particles through their anomalous specific ionization energy loss in the ATLAS pixel detector. In absence of an excess of events above the expected backgrounds, limits are set on stable bottom and top squark R-hadrons, see Table 5.
- 3 KHACHATRYAN 15AK looked in a data set corresponding to  $\text{fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$ , and a search interval corresponding to 281 h of trigger lifetime, for long-lived particles that have stopped in the CMS detector. No evidence for an excess over the expected background in a cloud interaction model is observed. Assuming the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  and lifetimes between  $1 \mu\text{s}$  and  $1000 \text{ s}$ , limits are derived on  $\tilde{t}$  production as a function of  $m_{\tilde{\chi}_1^0}$ , see Figs. 4 and 7. The exclusions require that  $m_{\tilde{\chi}_1^0}$  is kinematically consistent with the minimum values of the jet energy thresholds used.
- 4 AAD 13AA searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events containing colored long-lived particles that hadronize forming R-hadrons. No significant excess above the expected background was found. Long-lived R-hadrons containing a  $\tilde{t}$  are excluded for masses up to 683 GeV at 95% C.L. in a general interaction model. Also, limits independent of the fraction of R-hadrons that arrive charged in the muon system were derived, see Fig. 6.
- 5 AAD 13AA searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events containing colored long-lived particles that hadronize forming R-hadrons. No significant excess above the expected background was found. Long-lived R-hadrons containing a  $\tilde{b}$  are excluded for masses up to 612 GeV at 95% C.L. in a general interaction model. Also, limits independent of the fraction of R-hadrons that arrive charged in the muon system were derived, see Fig. 6.
- 6 AAD 13BC searched in  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  and in  $22.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for bottom squark R-hadrons that have come to rest within the ATLAS calorimeter and decay at some later time to hadronic jets and a neutralino. In absence of an excess of events above the expected backgrounds, limits are set on sbottom masses for the decay  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$ , for different lifetimes, and for a neutralino mass of 100 GeV, see their Table 6 and Fig 10.
- 7 AAD 13BC searched in  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  and in  $22.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for bottom squark R-hadrons that have come to rest within the ATLAS calorimeter and decay at some later time to hadronic jets and a neutralino. In absence of an excess of events above the expected backgrounds, limits are set on stop masses for the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$ , for different lifetimes, and for a neutralino mass of 100 GeV, see their Table 6 and Fig 10.

See key on page 601

## Searches Particle Listings Supersymmetric Particle Searches

<sup>8</sup> CHATRCHYAN 13AB looked in  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  and in  $18.8 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with heavy stable particles, identified by their anomalous  $dE/dx$  in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of  $\tilde{t}_1$ 's. No evidence for an excess over the expected background is observed. Limits are derived for pair production of stops as a function of mass in the cloud interaction model (see Fig. 8 and Table 6). In the charge-suppressed model, the limit decreases to 818 GeV.

### $\tilde{b}$ (Sbottom) MASS LIMIT

Limits in  $e^+e^-$  depend on the mixing angle of the mass eigenstate  $\tilde{b}_1 = \tilde{b}_L \cos\theta_b + \tilde{b}_R \sin\theta_b$ . Coupling to the  $Z$  vanishes for  $\theta_b \sim 1.17$ . As a consequence, no absolute constraint in the mass region  $\lesssim 40 \text{ GeV}$  is available in the literature at this time from  $e^+e^-$  collisions. In the Listings below, we use  $\Delta m = m_{\tilde{b}_1} - m_{\tilde{\chi}_1^0}$ .

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;600</b>	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{b} \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 250 \text{ GeV}$
<b>&gt;440</b>	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^{(*)}\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 60 \text{ GeV}, m_{\tilde{b}} - m_{\tilde{\chi}_1^\pm} < m_t$
none 300–650	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{b} \rightarrow \tilde{b}\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 60 \text{ GeV}, m_{\tilde{\chi}_2^0} > 250 \text{ GeV}$
<b>&gt;640</b>	95	<sup>2</sup> KHACHATRYAN...15AF	CMS	$\tilde{b} \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0$
<b>&gt;650</b>	95	<sup>3</sup> KHACHATRYAN...15AH	CMS	$\tilde{b} \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0$
<b>&gt;250</b>	95	<sup>3</sup> KHACHATRYAN...15AH	CMS	$\tilde{b} \rightarrow b\tilde{\chi}_1^0, m_{\tilde{b}} - m_{\tilde{\chi}_1^0} < 10 \text{ GeV}$
<b>&gt;570</b>	95	<sup>4</sup> KHACHATRYAN...15I	CMS	$\tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 50 \text{ GeV}, 150 < m_{\tilde{\chi}_1^\pm} < 300 \text{ GeV}$
<b>&gt;255</b>	95	<sup>5</sup> AAD	14T ATLS	$\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0, m_{\tilde{b}_1} - m_{\tilde{\chi}_1^0} \approx m_b$
<b>&gt;400</b>	95	<sup>6</sup> CHATRCHYAN 14AH	CMS	jets + $\cancel{E}_T, \tilde{b} \rightarrow b\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$
		<sup>7</sup> CHATRCHYAN 14R	CMS	$\geq 3e^\pm, \tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>8</sup> KHACHATRYAN...15AD	CMS	$\ell^\pm\ell^\mp + \text{jets} + \cancel{E}_T, \tilde{b} \rightarrow b\ell^\pm\ell^\mp\tilde{\chi}_1^0$
none 340–600	95	<sup>9</sup> AAD	14AX ATLS	$\geq 3 \text{ } b\text{-jets} + \cancel{E}_T, \tilde{b} \rightarrow b\tilde{\chi}_2^0$ simplified model with $\tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 60 \text{ GeV}, m_{\tilde{\chi}_2^0} = 300 \text{ GeV}$
<b>&gt;440</b>	95	<sup>10</sup> AAD	14E ATLS	$\ell^\pm\ell^\pm(\ell^\mp\ell^\mp) + \text{jets}, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$ with $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^\pm} = 2 m_{\tilde{\chi}_1^0}$
<b>&gt;500</b>	95	<sup>11</sup> CHATRCHYAN 14H	CMS	same-sign $\ell^\pm\ell^\pm, \tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^\pm} = 2 \text{ GeV}, m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
<b>&gt;620</b>	95	<sup>12</sup> AAD	13AU ATLS	$2 \text{ } b\text{-jets} + \cancel{E}_T, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 120 \text{ GeV}$
<b>&gt;550</b>	95	<sup>13</sup> CHATRCHYAN 13AT	CMS	jets + $\cancel{E}_T, \tilde{b} \rightarrow b\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$
<b>&gt;600</b>	95	<sup>14</sup> CHATRCHYAN 13T	CMS	jets + $\cancel{E}_T, \tilde{b} \rightarrow b\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 0 \text{ GeV}$
<b>&gt;450</b>	95	<sup>15</sup> CHATRCHYAN 13V	CMS	same-sign $\ell^\pm\ell^\pm + \geq 2 \text{ } b\text{-jets}, \tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$
<b>&gt;390</b>		<sup>16</sup> AAD	12AN ATLS	$\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^0} < 60 \text{ GeV}$
<b>&gt;410</b>	95	<sup>17</sup> CHATRCHYAN 12AI	CMS	$\ell^\pm\ell^\pm + b\text{-jets} + \cancel{E}_T$
		<sup>18</sup> CHATRCHYAN 12Bo	CMS	$\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$ , simplified model, $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$
<b>&gt;294</b>	95	<sup>19</sup> AAD	11K ATLS	stable $b$
		<sup>20</sup> AAD	11o ATLS	$\tilde{g} \rightarrow \tilde{b}_1 b, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 60 \text{ GeV}$
		<sup>21</sup> CHATRCHYAN 11D	CMS	$\tilde{b}, \tilde{t} \rightarrow b$
<b>&gt;230</b>	95	<sup>22</sup> AALTONEN	10R CDF	$\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 70 \text{ GeV}$
<b>&gt;247</b>	95	<sup>23</sup> ABAZOV	10L D0	$\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0 \text{ GeV}$

- <sup>1</sup> AAD 15CJ searched in  $20 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for evidence of third generation squarks by combining a large number of searches covering various final states. Limits on the sbottom mass are shown, either assuming the  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  decay, see Fig. 11, or assuming the  $\tilde{b} \rightarrow t\tilde{\chi}_1^\pm$  decay, with  $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\tilde{\chi}_1^0$ , see Fig. 12a, or assuming the  $\tilde{b} \rightarrow b\tilde{\chi}_2^0$  decay, with  $\tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0$ , see Fig. 12b. Interpretations in the pMSSM are also discussed, see Figures 13–15.
- <sup>2</sup> KHACHATRYAN 15AF searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with at least two energetic jets and significant  $\cancel{E}_T$ , using the transverse mass variable  $M_{T2}$  to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in simplified models where the decay  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 12. See also Table 5. Exclusions in the CMSSM, assuming  $\tan\beta = 30$ ,  $A_0 = -2 \max(m_0, m_{1/2})$  and  $\mu > 0$ , are also presented, see Fig. 15.
- <sup>3</sup> KHACHATRYAN 15AH searched in  $19.4$  or  $19.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events containing either a fully reconstructed top quark, or events containing dijets requiring one or both jets to originate from  $b$ -quarks, or events containing a mono-jet. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in simplified models where the decay  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 12. Limits are also set in a simplified model where the decay  $\tilde{b} \rightarrow c\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 12.
- <sup>4</sup> KHACHATRYAN 15I searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events in which  $b$ -jets and four  $W$ -bosons are produced. Five individual search channels are combined (fully hadronic, single lepton, same-sign dilepton, opposite-sign dilepton, multi-lepton). No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in a simplified model where the decay  $\tilde{b} \rightarrow t\tilde{\chi}_1^\pm$ , with  $\tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$ , takes place with a branching ratio of 100%, see Fig. 7.
- <sup>5</sup> AAD 14T searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for monojet-like events. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which assume that the decay  $\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$  takes place 100% of the time, see Fig. 12.
- <sup>6</sup> CHATRCHYAN 14AH searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with at least two energetic jets and significant  $\cancel{E}_T$ , using the razor variables ( $M_R$  and  $R_2$ ) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a  $b$ -quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , are also presented, see Fig. 26.
- <sup>7</sup> CHATRCHYAN 14R searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a simplified model where the decay  $\tilde{b} \rightarrow t\tilde{\chi}_1^\pm$ , with  $\tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$ , takes place with a branching ratio of 100%, see Fig. 11.
- <sup>8</sup> KHACHATRYAN 15AD searched in  $19.4 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two opposite-sign same flavor isolated leptons featuring either a kinematic edge, or a peak at the  $Z$ -boson mass, in the invariant mass spectrum. No evidence for a statistically significant excess over the expected SM backgrounds is observed and 95% C.L. exclusion limits are derived in a simplified model of sbottom pair production where the sbottom decays into a  $b$ -quark, two opposite-sign dileptons and a neutralino LSP, through an intermediate state containing either an off-shell  $Z$ -boson or a slepton, see Fig. 8.
- <sup>9</sup> AAD 14AX searched in  $20.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for the strong production of supersymmetric particles in events containing either zero or at least one high- $p_T$  lepton, large missing transverse momentum, high jet multiplicity and at least three jets identified as originating from  $b$ -quarks. No excess over the expected SM background is observed. Limits are derived in mSUGRA/CMSSM models with  $\tan\beta = 30$ ,  $A_0 = -2 m_0$  and  $\mu > 0$ , see their Fig. 14. Also, exclusion limits are set in simplified models containing scalar bottom quarks, where the decay  $\tilde{b} \rightarrow b\tilde{\chi}_2^0$  and  $\tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see their Figures 11.
- <sup>10</sup> AAD 14E searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for strongly produced supersymmetric particles in events containing jets and two same-sign leptons or three leptons. The search also utilises jets originating from  $b$ -quarks, missing transverse momentum and other variables. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing bottom, see Fig. 7. Limits are also derived in the mSUGRA/CMSSM, bRPV and GMSB models, see their Fig. 8.
- <sup>11</sup> CHATRCHYAN 14H searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in a simplified models where the decay  $\tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, with varying mass of the  $\tilde{\chi}_1^\pm$ , for  $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$ , see Fig. 6.
- <sup>12</sup> AAD 13AU searched in  $20.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events containing two jets identified as originating from  $b$ -quarks and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks. Assuming that the decay  $\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$  takes place 100% of the time, a  $\tilde{b}_1$  mass below 620 GeV is excluded for  $m_{\tilde{\chi}_1^0} < 120 \text{ GeV}$ . For more details, see their Fig. 5.
- <sup>13</sup> CHATRCHYAN 13AT provides interpretations of various searches for supersymmetry by the CMS experiment based on  $4.73\text{--}4.98 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  in the framework of simplified models. Limits are set on the sbottom mass in a simplified models where sbottom quarks are pair-produced and the decay  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 4.
- <sup>14</sup> CHATRCHYAN 13T searched in  $11.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with at least two energetic jets and significant  $\cancel{E}_T$ , using the  $\alpha_T$  variable to discriminate between processes with genuine and misreconstructed  $\cancel{E}_T$ . No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\tilde{b} \rightarrow b\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 8 and Table 9.

# Searches Particle Listings

## Supersymmetric Particle Searches

- <sup>15</sup> CHATRCHYAN 13v searched in  $10.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two isolated same-sign dileptons and at least two  $b$ -jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the bottom mass in a simplified models where the decay  $\tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm \tilde{\chi}_1^0$  takes place with a branching ratio of 100%, with varying mass of the  $\tilde{\chi}_1^\pm$ , for  $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$ , see Fig. 4.
- <sup>16</sup> AAD 12AN searched in  $2.05 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for scalar bottom quarks in events with large missing transverse momentum and two  $b$ -jets in the final state. The data are found to be consistent with the Standard Model expectations. Limits are set in an  $R$ -parity conserving minimal supersymmetric scenario, assuming  $B(\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0) = 100\%$ , see their Fig. 2.
- <sup>17</sup> CHATRCHYAN 12Ai looked in  $4.98 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with two same-sign leptons ( $e, \mu$ ), but not necessarily same flavor, at least 2  $b$ -jets and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in a simplified model for sbottom pair production, where the sbottom decays through  $\tilde{b}_1 \rightarrow t\tilde{\chi}_1 W$ , see Fig. 8.
- <sup>18</sup> CHATRCHYAN 12Bo searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for scalar bottom quarks in events with large missing transverse momentum and two  $b$ -jets in the final state. The data are found to be consistent with the Standard Model expectations. Limits are set in an  $R$ -parity conserving minimal supersymmetric scenario, assuming  $B(\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0) = 100\%$ , see their Fig. 2.
- <sup>19</sup> AAD 11K looked in  $34 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with heavy stable particles, identified by their anomalous  $dE/dx$  in the tracker or time of flight in the tile calorimeter, from pair production of  $\tilde{b}$ . No evidence for an excess over the SM expectation is observed and limits on the mass are derived for pair production of sbottom, see Fig. 4.
- <sup>20</sup> AAD 11o looked in  $35 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with jets, of which at least one is a  $b$ -jet, and  $\cancel{E}_T$ . No excess above the Standard Model was found. Limits are derived in the  $(m_{\tilde{g}}, m_{\tilde{b}_1})$  plane (see Fig. 2) under the assumption of 100% branching ratios and  $\tilde{b}_1$  being the lightest squark. The quoted limit is valid for  $m_{\tilde{b}_1} < 500 \text{ GeV}$ . A similar approach for  $\tilde{t}_1$  as the lightest squark with  $\tilde{g} \rightarrow \tilde{t}_1 t$  and  $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$  with 100% branching ratios leads to a gluino mass limit of  $520 \text{ GeV}$  for  $130 < m_{\tilde{t}_1} < 300 \text{ GeV}$ . Limits are also derived in the CMSSM  $(m_0, m_{1/2})$  plane for  $\tan\beta = 40$ , see Fig. 4, and in scenarios based on the gauge group  $SO(10)$ .
- <sup>21</sup> CHATRCHYAN 11D looked in  $35 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with  $\geq 2$  jets, at least one of which is  $b$ -tagged, and  $\cancel{E}_T$ , where the  $b$ -jets are decay products of  $\tilde{t}$  or  $\tilde{b}$ . No evidence for an excess over the expected background is observed. Limits are derived in the CMSSM  $(m_0, m_{1/2})$  plane for  $\tan\beta = 50$  (see Fig. 2).
- <sup>22</sup> AALTONEN 10R searched in  $2.65 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$  for events with  $\cancel{E}_T$  and exactly two jets, at least one of which is  $b$ -tagged. The results are in agreement with the SM prediction, and a limit on the cross section of  $0.1 \text{ pb}$  is obtained for the range of masses  $80 < m_{\tilde{b}_1} < 280 \text{ GeV}$  assuming that the sbottom decays exclusively to  $b\tilde{\chi}_1^0$ . The excluded mass region in the framework of conserved  $R_p$  is shown in a plane of  $(m_{\tilde{b}_1}, m_{\tilde{\chi}_1^0})$ , see their Fig. 2.
- <sup>23</sup> ABZOV 10L looked in  $5.2 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$  for events with at least 2  $b$ -jets and  $\cancel{E}_T$  from the production of  $\tilde{b}_1 \tilde{b}_1$ . No evidence for an excess over the SM expectation is observed, and a limit on the cross section is derived under the assumption of 100% branching ratio. The excluded mass region in the framework of conserved  $R_p$  is shown in a plane of  $(m_{\tilde{b}_1}, m_{\tilde{\chi}_1^0})$ , see their Fig. 3b. The exclusion also extends to  $m_{\tilde{\chi}_1^0} = 110 \text{ GeV}$  for  $160 < m_{\tilde{b}_1} < 200 \text{ GeV}$ .

### $\tilde{t}$ (Stop) MASS LIMIT

Limits depend on the decay mode. In  $e^+e^-$  collisions they also depend on the mixing angle of the mass eigenstate  $\tilde{t}_1 = \tilde{t}_L \cos\theta_t + \tilde{t}_R \sin\theta_t$ . The coupling to the  $Z$  vanishes when  $\theta_t = 0.98$ . In the Listings below, we use  $\Delta m \equiv m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0}$  or  $\Delta m \equiv m_{\tilde{t}_1} - m_{\tilde{\nu}}$ , depending on relevant decay mode. See also bounds in “ $\tilde{a}$  (Squark) MASS LIMIT.”

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>250	95	<sup>1</sup> AAD	15CJ ATLS	$B(\tilde{t} \rightarrow c\tilde{\chi}_1^0) + B(\tilde{t} \rightarrow bff'\tilde{\chi}_1^0) = 1, m_{\tilde{t}} - m_{\tilde{\chi}_1^0} = 10 \text{ GeV}$
>270	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{t} \rightarrow c\tilde{\chi}_1^0, m_{\tilde{t}} - m_{\tilde{\chi}_1^0} = 80 \text{ GeV}$
none, 200–700	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{t} \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0$
>500	95	<sup>1</sup> AAD	15CJ ATLS	$B(\tilde{t} \rightarrow t\tilde{\chi}_1^0) + B(\tilde{t} \rightarrow b\tilde{\chi}_1^\pm) = 1, \tilde{\chi}_1^\pm \rightarrow W^{(*)}\tilde{\chi}_1^0, m_{\tilde{\chi}_1^\pm} = 2m_{\tilde{\chi}_1^0}, m_{\tilde{\chi}_1^0} < 160 \text{ GeV}$
>600	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{t}_2 \rightarrow Z\tilde{t}_1, m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} = 180 \text{ GeV}, m_{\tilde{\chi}_1^0} = 0$
>600	95	<sup>1</sup> AAD	15CJ ATLS	$\tilde{t}_2 \rightarrow h\tilde{t}_1, m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} = 180 \text{ GeV}, m_{\tilde{\chi}_1^0} = 0$
none, 172.5–191	95	<sup>2</sup> AAD	15J ATLS	$\tilde{t} \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 1 \text{ GeV}$
>450	95	<sup>3</sup> KHACHATRY...15AF	CMS	$\tilde{t} \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0, m_{\tilde{t}} > m_t + m_{\tilde{\chi}_1^0}$

>560	95	<sup>4</sup> KHACHATRY...15AH	CMS	$\tilde{t} \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0, m_{\tilde{t}} > m_t + m_{\tilde{\chi}_1^0}$
>250	95	<sup>5</sup> KHACHATRY...15AH	CMS	$\tilde{t} \rightarrow c\tilde{\chi}_1^0, m_{\tilde{t}} - m_{\tilde{\chi}_1^0} < 10 \text{ GeV}$
none, 200–350	95	<sup>6</sup> KHACHATRY...15L	CMS	$\tilde{t} \rightarrow qq, R, \lambda_{312}'' \neq 0$
none, 200–385	95	<sup>6</sup> KHACHATRY...15L	CMS	$\tilde{t} \rightarrow qb, R, \lambda_{323}'' \neq 0$
<b>&gt;730</b>	95	<sup>7</sup> KHACHATRY...15X	CMS	$\tilde{t} \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 100 \text{ GeV}, m_{\tilde{t}} > m_t + m_{\tilde{\chi}_1^0}$
none 400–645	95	<sup>7</sup> KHACHATRY...15x	CMS	$\tilde{t} \rightarrow t\tilde{\chi}_1^0 \text{ or } \tilde{t} \rightarrow b\tilde{\chi}_1^\pm, m_{\tilde{\chi}_1^0} = 100 \text{ GeV}, m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} = 5 \text{ GeV}$
none 270–645	95	<sup>8</sup> AAD	14AJ ATLS	$\geq 4 \text{ jets} + \cancel{E}_T, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 30 \text{ GeV}$
none 250–550	95	<sup>8</sup> AAD	14AJ ATLS	$\geq 4 \text{ jets} + \cancel{E}_T, B(\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm) = 50\%, m_{\tilde{\chi}_1^\pm} = 2m_{\tilde{\chi}_1^0}, m_{\tilde{\chi}_1^0} < 60 \text{ GeV}$
none 210–640	95	<sup>9</sup> AAD	14BD ATLS	$\ell^\pm + \text{jets} + \cancel{E}_T, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 0 \text{ GeV}$
<b>&gt;500</b>	95	<sup>9</sup> AAD	14BD ATLS	$\ell^\pm + \text{jets} + \cancel{E}_T, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm, m_{\tilde{\chi}_1^\pm} = 2m_{\tilde{\chi}_1^0}, 100 \text{ GeV} < m_{\tilde{\chi}_1^0} < 150 \text{ GeV}$
none 150–445	95	<sup>10</sup> AAD	14F ATLS	$\ell^\pm \ell^\mp$ final state, $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm, m_{\tilde{t}_1} - m_{\tilde{\chi}_1^\pm} = 10 \text{ GeV}, m_{\tilde{\chi}_1^0} = 1 \text{ GeV}$
none 215–530	95	<sup>10</sup> AAD	14F ATLS	$\ell^\pm \ell^\mp$ final state, $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 1 \text{ GeV}$
>270	95	<sup>11</sup> AAD	14T ATLS	$\tilde{t}_1 \rightarrow c\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 200 \text{ GeV}$
<b>&gt;240</b>	95	<sup>11</sup> AAD	14T ATLS	$\tilde{t}_1 \rightarrow c\tilde{\chi}_1^0, m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} < 85 \text{ GeV}$
>255	95	<sup>11</sup> AAD	14T ATLS	$\tilde{t}_1 \rightarrow bff'\tilde{\chi}_1^0, m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0} \approx m_b$
>400	95	<sup>12</sup> CHATRCHYAN 14AH	CMS	$\text{jets} + \cancel{E}_T, \tilde{t} \rightarrow t\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$
		<sup>13</sup> CHATRCHYAN 14R	CMS	$\geq 3\ell^\pm, \tilde{t} \rightarrow (b\tilde{\chi}_1^\pm/t\tilde{\chi}_1^0), \tilde{\chi}_1^\pm \rightarrow (qq'/\ell\nu)\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow (H/Z)\tilde{G}, \text{GMSB, natural higgsino NLSP scenario}$
>740	95	<sup>14</sup> KHACHATRY...14T	CMS	$\tau + b\text{-jets}, R, LQ\bar{D}, \lambda_{333}'' \neq 0, \tilde{t} \rightarrow \tau b$ simplified model
>580	95	<sup>14</sup> KHACHATRY...14T	CMS	$\tau + b\text{-jets}, R, LQ\bar{D}, \lambda_{3jk}'' \neq 0 (j \neq 3), \tilde{t} \rightarrow \tilde{\chi}^\pm b, \tilde{\chi}^\pm \rightarrow q\bar{q}\tau^\pm$ simplified model
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>790	95	<sup>15</sup> KHACHATRY...15E	CMS	$\tilde{t}_1 \rightarrow b\ell, \text{RPV}, c\tau = 2 \text{ cm}$
>230		ROLBIECKI 15	THEO	$WW$ xsection, $\tilde{t}_1 \rightarrow bW\tilde{\chi}_1^0, m_{\tilde{t}_1} \simeq m_b + m_W + m_{\tilde{\chi}_1^0}$
>600	95	<sup>16</sup> AAD	14B ATLS	$Z + b\cancel{E}_T, \tilde{t}_2 \rightarrow Z\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 200 \text{ GeV}$
>540	95	<sup>16</sup> AAD	14B ATLS	$Z + b\cancel{E}_T, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}, \text{natural GMSB}, 100 \text{ GeV} < m_{\tilde{\chi}_1^0} < m_{\tilde{t}_1} - 10 \text{ GeV}$
>360	95	<sup>17</sup> CHATRCHYAN 14U	CMS	$\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow f'f'\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow H\tilde{G}$ simplified model, $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} = 5 \text{ GeV}, \text{GMSB}$
>215	95	CZAKON 14		$\tilde{t} \rightarrow t\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 10 \text{ GeV}$
		<sup>18</sup> KHACHATRY...14C	CMS	$\tilde{t}_2 \rightarrow H\tilde{t}_1 \text{ or } \tilde{t}_2 \rightarrow Z\tilde{t}_1$ simplified model

- <sup>1</sup> AAD 15CJ searched in  $20 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for evidence of third generation squarks by combining a large number of searches covering various final states. Stop decays with and without charginos in the decay chain are considered and summaries of all ATLAS Run 1 searches for direct stop production can be found in Fig. 4 (no intermediate charginos) and Fig. 7 (intermediate charginos). Limits are set on stop masses in compressed mass regions regions, with  $B(\tilde{t} \rightarrow c\tilde{\chi}_1^0) + B(\tilde{t} \rightarrow bff'\tilde{\chi}_1^0) = 1$ , see Fig. 5. Limits are also set on stop masses assuming that both the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  and  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$  are possible, with both their branching ratios summing up to 1, assuming  $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\tilde{\chi}_1^0$  and  $m_{\tilde{\chi}_1^\pm} = 2m_{\tilde{\chi}_1^0}$ , see Fig. 6. Limits on the mass of the next-to-lightest stop  $\tilde{t}_2$ , decaying either to  $Z\tilde{t}_1, h\tilde{t}_1$  or  $t\tilde{\chi}_1^0$ , are also presented, see Figs. 9 and 10. Interpretations in the pMSSM are also discussed, see Figs 13–15.
- <sup>2</sup> AAD 15J interpreted the measurement of spin correlations in  $t\bar{t}$  production using  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  in exclusion limits on the pair production of light  $\tilde{t}_1$  squarks with masses similar to the top quark mass. The  $\tilde{t}_1$  is assumed to decay through  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$  with predominantly right-handed top and a 100% branching ratio. The data are found to be consistent with the Standard Model expectations and masses between the top quark mass and 191 GeV are excluded, see their Fig. 2.



- <sup>3</sup> KHACHATRYAN 15AF searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least two energetic jets and significant  $E_T$ , using the transverse mass variable  $M_{T2}$  to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 12. See also Table 5. Exclusions in the CMSSM, assuming  $\tan\beta = 30$ ,  $A_0 = -2 \max(m_0, m_{1/2})$  and  $\mu > 0$ , are also presented, see Fig. 15.
- <sup>4</sup> KHACHATRYAN 15AH searched in 19.4 or 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing either a fully reconstructed top quark, or events containing dijets requiring one or both jets to originate from  $b$ -quarks, or events containing a mono-jet. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 9. Limits are also set in simplified models where the decays  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  and  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ , with  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} = 5$  GeV, each take place with a branching ratio of 50%, see Fig. 10, or with other fractions, see Fig. 11. Finally, limits are set in a simplified model where the decay  $\tilde{t} \rightarrow c\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Figs. 9, 10 and 11.
- <sup>5</sup> KHACHATRYAN 15AH searched in 19.4 or 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing either a fully reconstructed top quark, or events containing dijets requiring one or both jets to originate from  $b$ -quarks, or events containing a mono-jet. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 9. Limits are also set in simplified models where the decays  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  and  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ , with  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} = 5$  GeV, each take place with a branching ratio of 50%, see Fig. 10, or with other fractions, see Fig. 11. Finally, limits are set in a simplified model where the decay  $\tilde{t} \rightarrow c\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Figs. 9, 10, and 11.
- <sup>6</sup> KHACHATRYAN 15L searched in 19.4 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for pair production of heavy resonances decaying to pairs of jets in four jet events. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in  $R$ -parity-violating supersymmetry models where  $\tilde{t} \rightarrow q\bar{q}$  ( $\lambda_{312} \neq 0$ ), see Fig. 6 (top) and  $\tilde{t} \rightarrow q\bar{b}$  ( $\lambda_{323} \neq 0$ ), see Fig. 6 (bottom).
- <sup>7</sup> KHACHATRYAN 15x searched in 19.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least two energetic jets, at least one of which is required to originate from a  $b$  quark, possibly a lepton, and significant  $E_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  and the decay  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ , with  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} = 5$  GeV, take place with branching ratios varying between 0 and 100%, see Figs. 15, 16 and 17.
- <sup>8</sup> AAD 14AJ searched in 20.1 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing four or more jets and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which either assume that the decay  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$  takes place 100% of the time, see Fig. 8, or that this decay takes place 50% of the time, while the decay  $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$  takes place the other 50% of the time, see Fig. 9.
- <sup>9</sup> AAD 14BD searched in 20 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing one isolated lepton, jets and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which either assume that the decay  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$  takes place 100% of the time, see Fig. 15, or the decay  $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$  takes place 100% of the time, see Fig. 16–22. For the mixed decay scenario, see Fig. 23.
- <sup>10</sup> AAD 14F searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing two leptons ( $e$  or  $\mu$ ), and possibly jets and missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which either assume that the decay  $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$  takes place 100% of the time, see Figs. 14–17 and 20, or that the decay  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$  takes place 100% of the time, see Figs. 18 and 19.
- <sup>11</sup> AAD 14T searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for monojet-like and  $c$ -tagged events. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which assume that the decay  $\tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$  takes place 100% of the time, see Fig. 9 and 10. The results of the monojet-like analysis are also interpreted in terms of stop pair production in the four-body decay  $\tilde{t}_1 \rightarrow b\bar{t}\bar{t}'\tilde{\chi}_1^0$ , see Fig. 11.
- <sup>12</sup> CHATRCHYAN 14AH searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with at least two energetic jets and significant  $E_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a  $b$ -quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\tilde{t} \rightarrow t\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , are also presented, see Fig. 26.
- <sup>13</sup> CHATRCHYAN 14R searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in a natural higgsino NLSP simplified model (GMSB) where the decay  $\tilde{t} \rightarrow b\tilde{\chi}_1^\pm$ , with  $\tilde{\chi}_1^\pm \rightarrow (qq'/\ell\nu)H$ ,  $Z\bar{G}$ , takes place with a branching ratio of 100% (the particles between brackets have a soft  $p_T$  spectrum), see Figs. 4–6.

- <sup>14</sup> KHACHATRYAN 14T searched in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with  $\tau$ -leptons and  $b$ -quark jets, possibly with extra light-flavour jets. No excess above the Standard Model expectations is observed. Limits are set on stop masses in  $R$  SUSY models with  $L\bar{Q}\bar{D}$  couplings, in two simplified models. In the first model, the decay  $\tilde{t} \rightarrow \tau b$  is considered, with  $\lambda_{333} \neq 0$ , see Fig. 3. In the second model, the decay  $\tilde{t} \rightarrow \tilde{\chi}^\pm b$ , with the subsequent decay  $\tilde{\chi}^\pm \rightarrow qq\tau^\pm$  is considered, with  $\lambda'_{3jk} \neq 0$  and the mass splitting between the top squark and the charging chosen to be 100 GeV, see Fig. 4.
- <sup>15</sup> KHACHATRYAN 15E searched for long-lived particles decaying to leptons in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV. Events were selected with an electron and muon with opposite charges and each with transverse impact parameter values between 0.02 and 2 cm. Limits are set on SUSY benchmark models with pair production of top squarks decaying into an  $e\mu$  final state via RPV interactions. See their Fig. 2.
- <sup>16</sup> AAD 14B searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing a  $Z$  boson, with or without additional leptons, plus jets originating from  $b$ -quarks and significant missing transverse momentum. No excess over the expected SM background is observed. Limits are derived in simplified models featuring  $\tilde{t}_2$  production, with  $\tilde{t}_2 \rightarrow Z\tilde{t}_1$ ,  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$  with a 100% branching ratio, see Fig. 4, and in the framework of natural GMSB, see Fig. 6.
- <sup>17</sup> CHATRCHYAN 14U searched in 19.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for evidence of direct pair production of top squarks, with Higgs bosons in the decay chain. The search is performed using a selection of events containing two Higgs bosons, each decaying to a photon pair, missing transverse energy and possibly  $b$ -quark jets. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of a “natural SUSY” simplified model where the decays  $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$ , with  $\tilde{\chi}_1^\pm \rightarrow f\bar{f}'\tilde{\chi}_1^0$ , and  $\tilde{\chi}_1^0 \rightarrow H\bar{G}$ , all happen with 100% branching ratio, see Fig. 4.
- <sup>18</sup> KHACHATRYAN 14C searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for evidence of direct pair production of top squarks, with Higgs or  $Z$ -bosons in the decay chain. The search is performed using a selection of events containing leptons and  $b$ -quark jets. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of a simplified model with pair production of a heavier top-squark mass eigenstate  $\tilde{t}_2$  decaying to a lighter top-squark eigenstate  $\tilde{t}_1$  via either  $\tilde{t}_2 \rightarrow H\tilde{t}_1$  or  $\tilde{t}_2 \rightarrow Z\tilde{t}_1$ , followed in both cases by  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$ . The interpretation is performed in the region where the mass difference between the  $\tilde{t}_1$  and  $\tilde{\chi}_1^0$  is approximately equal to the top-quark mass, which is not probed by searches for direct  $\tilde{t}_1$  pair production, see Figs. 5 and 6. The analysis excludes top squarks with masses  $m_{\tilde{t}_2} < 575$  GeV and  $m_{\tilde{t}_1} < 400$  GeV at 95% C.L.

Heavy  $\tilde{g}$  (Gluino) MASS LIMIT

For  $m_{\tilde{g}} > 60$ –70 GeV, it is expected that gluinos would undergo a cascade decay via a number of neutralinos and/or charginos rather than undergo a direct decay to photinos as assumed by some papers. Limits obtained when direct decay is assumed are usually higher than limits when cascade decays are included.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	C.L.	DOCUMENT ID	TECN	COMMENT
> 820	95	1 AAD	15BG ATLS	GGM, $\tilde{g} \rightarrow q\bar{q}Z\bar{G}$ , $\tan\beta = 30$ , $\mu > 600$ GeV
> 850	95	1 AAD	15BG ATLS	GGM, $\tilde{g} \rightarrow q\bar{q}Z\bar{G}$ , $\tan\beta = 1.5$ , $\mu > 450$ GeV
>1150	95	2 AAD	15BV ATLS	general RPC $\tilde{g}$ decays, $m_{\tilde{\chi}_1^0} < 100$ GeV
> 700	95	3 AAD	15BX ATLS	$\tilde{g} \rightarrow X_{\tilde{\chi}_1^0}$ , independent of $m_{\tilde{\chi}_1^0}$
>1290	95	4 AAD	15CA ATLS	$\geq 2\gamma + E_T$ , GGM, bino-like NLSP, any NLSP mass
>1260	95	4 AAD	15CA ATLS	$\geq 1\gamma + b$ -jets + $E_T$ , GGM, higgsino-bino admix. NLSP and $\mu < 0$ , $m(\text{NLSP}) > 450$ GeV
>1140	95	4 AAD	15CA ATLS	$\geq 1\gamma + \text{jets} + E_T$ , GGM, higgsino-bino admixture NLSP, all $\mu > 0$
>1225	95	5 KHACHATRY...15AF	CMS	$\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 0$
>1300	95	5 KHACHATRY...15AF	CMS	$\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 0$
>1225	95	5 KHACHATRY...15AF	CMS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 0$
>1550	95	5 KHACHATRY...15AF	CMS	CMSSM, $\tan\beta=30$ , $m_{\tilde{g}}=m_{\tilde{q}}$ , $A_0=-2\max(m_0, m_{1/2})$ , $\mu > 0$
>1150	95	5 KHACHATRY...15AF	CMS	CMSSM, $\tan\beta=30$ , $A_0=-2\max(m_0, m_{1/2})$ , $\mu > 0$
>1280	95	6 KHACHATRY...15I	CMS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 0$
>1310	95	7 KHACHATRY...15X	CMS	$\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 100$ GeV
>1175	95	7 KHACHATRY...15X	CMS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 100$ GeV
>1330	95	8 AAD	14AE ATLS	jets + $E_T$ , $\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 0$ GeV
>1700	95	8 AAD	14AE ATLS	jets + $E_T$ , mSUGRA/CMSSM, $m_{\tilde{q}} = m_{\tilde{g}}$
>1090	95	9 AAD	14AG ATLS	$\tau$ + jets + $E_T$ , natural Gauge Mediation
>1600	95	9 AAD	14AG ATLS	$\tau$ + jets + $E_T$ , mGMSB, $M_{\text{mess}} = 250$ GeV, $N_5 = 3$ , $\mu > 0$ , $C_{\text{grav}} = 1$
>1350	95	10 AAD	14X ATLS	$\geq 4\ell^\pm$ , $\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$ , $\tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \nu$ , $R$

# Searches Particle Listings

## Supersymmetric Particle Searches

> 640	95	11 AAD	14x ATLS	$\geq 4\ell^\pm, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell^\pm \ell^\mp \tilde{G}, \tan\beta = 30$ , GGM	>1400	95	22 AAD	15CB ATLS	jets or $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$ , Split SUSY, $m_{\tilde{\chi}_1^0} = 100$ GeV and
>1000	95	12 CHATRCHYAN14AH CMS		jets + $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50$ GeV	>1500	95	22 AAD	15CB ATLS	$15 < c\tau < 300$ mm $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$ , Split SUSY, $m_{\tilde{\chi}_1^0} = 100$ GeV and $20 < c\tau < 250$ mm
>1350	95	12 CHATRCHYAN14AH CMS		jets + $\cancel{E}_T$ , CMSSM, $m_{\tilde{g}} = m_{\tilde{q}}$	>1000	95	23 AAD	15x ATLS	$\geq 10$ jets, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}q$ (RPV), $m_{\tilde{\chi}_1^0} = 500$ GeV
>1000	95	13 CHATRCHYAN14AH CMS		jets + $\cancel{E}_T, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 50$ GeV	> 917	95	23 AAD	15x ATLS	$\geq 6,7$ jets, $\tilde{g} \rightarrow q\tilde{q}q$ , (light-quark, $\lambda''$ couplings, RPV)
>1160	95	15 CHATRCHYAN14I CMS		ets + $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 100$ GeV	> 929	95	23 AAD	15x ATLS	$\geq 6,7$ jets, $\tilde{g} \rightarrow q\tilde{q}q$ , (b-quark, $\lambda''$ couplings, RPV)
>1130	95	15 CHATRCHYAN14I CMS		multijets + $\cancel{E}_T, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 100$ GeV	>1300	95	24 KHACHATRY...15AD CMS		$\ell^\pm \ell^\mp + \text{jets} + \cancel{E}_T$ , GMSB, $\tilde{g} \rightarrow q\tilde{q}Z\tilde{G}$
>1210	95	15 CHATRCHYAN14I CMS		multijets + $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}W/Z\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 100$ GeV	> 800	95	25 KHACHATRY...15AZ CMS		$\geq 2\gamma, \geq 1$ jet, (Razor), binolike NLSP, $m_{\tilde{\chi}_1^0} = 375$ GeV
>1260	95	16 CHATRCHYAN14N CMS		$1\ell^\pm + \text{jets} + \geq 2b\text{-jets}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} = 0$ GeV, $m_{\tilde{t}} > m_{\tilde{g}}$	>1280	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T$ , CMSSM
> 650	95	17 CHATRCHYAN14P CMS		$\tilde{g} \rightarrow jjj, R$	>1250	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T, \tilde{g} \rightarrow b_1\tilde{b}\tilde{\chi}_1^0$ simplified model, $b_1 \rightarrow b\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 60$ GeV, $m_{b_1} < 900$ GeV
none 200–835	95	17 CHATRCHYAN14P CMS		$\tilde{g} \rightarrow bj\bar{j}, R$	>1190	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T, \tilde{g} \rightarrow \tilde{t}_1 t\tilde{\chi}_1^0$ simplified model, $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 60$ GeV, $m_{\tilde{t}_1} < 1000$ GeV
		18 CHATRCHYAN14R CMS		$\geq 3\ell^\pm, (\tilde{g}/\tilde{q}) \rightarrow q\ell^\pm \ell^\mp \tilde{G}$ simplified model, GMSB, slepton co-NLSP sceario	>1180	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T, \tilde{g} \rightarrow \tilde{t}_1 t\tilde{\chi}_1^0$ simplified model, $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$ , $m_{\tilde{\chi}_1^\pm} = 2m_{\tilde{\chi}_1^0}$ , $m_{\tilde{\chi}_1^0} = 60$ GeV, $m_{\tilde{t}_1} < 1000$ GeV
		19 CHATRCHYAN14R CMS		$\geq 3\ell^\pm, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$ simplified model	>1250	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 400$ GeV
	95	20 AAD	15AB ATLS	$\tilde{g} \rightarrow \tilde{S}g, c\tau = 1$ m, $\tilde{S} \rightarrow \tilde{S}\tilde{G}$ and $\tilde{S} \rightarrow g\tilde{g}$ , BR = 100%	>1340	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 400$ GeV
>1600	95	21 AAD	15AI ATLS	$\ell^\pm + \text{jets} + \cancel{E}_T$	>1300	95	26 AAD	14AX ATLS	$\geq 3$ b-jets + $\cancel{E}_T, \tilde{g} \rightarrow t\tilde{b}\tilde{\chi}_1^\pm$ simplified model, $\tilde{\chi}_1^\pm \rightarrow f'\ell'\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_1^0} = 2$ GeV, $m_{\tilde{\chi}_1^0} < 300$ GeV
>1280	95	2 AAD	15BV ATLS	pMSSM, $M_1 = 60$ GeV, $m_{\tilde{q}} < 1500$ GeV	> 950	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$ simplified model
>1100	95	2 AAD	15BV ATLS	mSUGRA, $m_0 > 2$ TeV	>1000	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow t\tilde{t}_1$ with $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$ simplified model, $m_{\tilde{t}_1} < 200$ GeV, $m_{\tilde{\chi}_1^\pm} = 118$ GeV, $m_{\tilde{\chi}_1^0} = 60$ GeV
>1330	95	2 AAD	15BV ATLS	via $\tilde{\tau}$ , natural GMSB, all $m_{\tilde{\tau}}$	> 640	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow t\tilde{t}_1$ with $\tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$ simplified model, $m_{\tilde{t}_1} = m_{\tilde{\chi}_1^0} + 20$ GeV
>1500	95	2 AAD	15BV ATLS	jets + $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ , $m_{\tilde{\chi}_1^0} = 1$ GeV	> 850	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow t\tilde{t}_1$ with $\tilde{t}_1 \rightarrow b\tilde{s}$ simplified model, $R$
>1650	95	2 AAD	15BV ATLS	jets + $\cancel{E}_T, m_{\tilde{g}} = m_{\tilde{q}}, m_{\tilde{\chi}_1^0} = 1$ GeV	> 860	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow q\tilde{q}'\tilde{\chi}_1^\pm$ , $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\pm\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^\pm} = 2$ $m_{\tilde{\chi}_1^0}$ , $m_{\tilde{\chi}_1^0} < 400$ GeV
> 850	95	2 AAD	15BV ATLS	jets + $\cancel{E}_T, \tilde{g} \rightarrow g\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 550$ GeV	>1040	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow q\tilde{q}'\tilde{\chi}_1^\pm$ , $\tilde{\chi}_1^\pm \rightarrow W^{(*)}\pm\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z^{(*)}\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^0} < 520$ GeV
>1270	95	2 AAD	15BV ATLS	jets + $\cancel{E}_T, \tilde{g} \rightarrow q\tilde{q}W\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 100$ GeV	>1200	95	27 AAD	14E ATLS	$\ell^\pm \ell^\pm (\ell^\mp) + \text{jets}, \tilde{g} \rightarrow q\tilde{q}'\tilde{\chi}_1^\pm/\tilde{\chi}_2^0, \tilde{\chi}_1^\pm \rightarrow \ell^\pm \nu\tilde{\chi}_1^0, \tilde{\chi}_2^0 \rightarrow \ell^\pm \ell^\mp (\nu\nu)\tilde{\chi}_1^0$ simplified model
>1150	95	2 AAD	15BV ATLS	jets + $\ell^\pm \ell^\pm, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 100$ GeV	>1050	95	28 CHATRCHYAN14H CMS		same-sign $\ell^\pm \ell^\pm, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$ simplified model, massless $\tilde{\chi}_1^0$
>1320	95	2 AAD	15BV ATLS	jets + $\ell^\pm \ell^\pm, \tilde{g}$ decays via sleptons, $m_{\tilde{\chi}_1^0} = 100$ GeV	> 900	95	29 CHATRCHYAN14H CMS		same-sign $\ell^\pm \ell^\pm, \tilde{g} \rightarrow q\tilde{q}'\tilde{\chi}_1^\pm$ , $\tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$ simplified model, $m_{\tilde{\chi}_1^\pm} = 0.5$ $m_{\tilde{g}}$ , massless $\tilde{\chi}_1^0$
>1220	95	2 AAD	15BV ATLS	$\tau, \tilde{q}$ decays via staus, $m_{\tilde{\chi}_1^0} = 100$ GeV					
>1310	95	2 AAD	15BV ATLS	b-jets, $\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 400$ GeV					
>1220	95	2 AAD	15BV ATLS	b-jets, $\tilde{g} \rightarrow \tilde{t}_1 t$ and $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0, m_{\tilde{t}_1} < 1000$ GeV					
>1180	95	2 AAD	15BV ATLS	b-jets, $\tilde{g} \rightarrow \tilde{t}_1 t$ and $\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm, m_{\tilde{t}_1} < 1000$ GeV, $m_{\tilde{\chi}_1^0} = 60$ GeV					
>1260	95	2 AAD	15BV ATLS	b-jets, $\tilde{g} \rightarrow \tilde{t}_1 t$ and $\tilde{g} \rightarrow c\tilde{\chi}_1^0$					
> 880	95	2 AAD	15BV ATLS	jets, $\tilde{g} \rightarrow \tilde{t}_1 t$ and $\tilde{t}_1 \rightarrow sb$ , RPV, $400 < m_{\tilde{t}_1} < 1000$ GeV					
>1200	95	2 AAD	15BV ATLS	b-jets, $\tilde{g} \rightarrow \tilde{b}_1 b$ and $\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0, m_{\tilde{b}_1} < 1000$ GeV					
>1250	95	2 AAD	15BV ATLS	b-jets, $\tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} < 400$ GeV					
none, 750–1250	95	2 AAD	15BV ATLS	b-jets, $\tilde{g}$ decay via offshell $\tilde{t}_1$ and $\tilde{b}_1, m_{\tilde{\chi}_1^0} < 500$ GeV					
> 600	95	22 AAD	15CB ATLS	$\ell, \tilde{g} \rightarrow (e/\mu)q\tilde{q}$ , RPV, benchmark gluino, neutralino masses					
	95	22 AAD	15CB ATLS	$\ell\ell/Z, \tilde{g} \rightarrow (e/\mu\mu/e\mu)q\tilde{q}$ , RPV, $m_{\tilde{\chi}_1^0} = 400$ GeV and $0.7 < c\tau_{\tilde{\chi}_1^0} < 3 \times 10^5$ mm					
>1100	95	22 AAD	15CB ATLS	jets, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$ , GGM, $m_{\tilde{\chi}_1^0} = 400$ GeV and $3 < c\tau_{\tilde{\chi}_1^0} < 500$ mm					

>1050	95	<sup>30</sup> CHATRCHYAN 14H	CMS	same-sign $e^\pm e^\pm, \bar{g} \rightarrow b\bar{\chi}_1^\pm$ , $\bar{\chi}_1^\pm \rightarrow W^\pm \bar{\chi}_1^0$ simplified model, $m_{\bar{\chi}_1^\pm} = 300$ GeV, $m_{\bar{\chi}_1^0}$ $= 50$ GeV
> 900	95	<sup>31</sup> CHATRCHYAN 14H		same-sign $e^\pm e^\pm, \bar{g} \rightarrow t\bar{b}$ sim- plified model, $\bar{g}$

- <sup>1</sup> AAD 15BG searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with jets, missing  $E_T$ , and two opposite-sign same flavor isolated leptons featuring either a kinematic edge, or a peak at the Z-boson mass, in the invariant mass spectrum. No evidence for a statistically significant excess over the expected SM backgrounds is observed and 95% C.L. exclusion limits are derived in a GGM simplified model of gluino pair production where the gluino decays into quarks, a Z-boson, and a massless gravitino LSP, see Fig. 12. Also, limits are set in simplified models with slepton/sneutrino intermediate states, see Fig. 13.
- <sup>2</sup> AAD 15bv summarized and extended ATLAS searches for gluinos and first- and second-generation squarks in final states containing jets and missing transverse momentum, with or without leptons or  $b$ -jets in the  $\sqrt{s} = 8$  TeV data set collected in 2012. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on the gluino mass in several R-parity conserving models, leading to a generalized constraint on gluino masses exceeding 1150 GeV for lightest supersymmetric particle masses below 100 GeV. See their Figs. 10, 19, 20, 21, 23, 25, 26, 29–37.
- <sup>3</sup> AAD 15bx interpreted the results of a wide range of ATLAS direct searches for supersymmetry, during the first run of the LHC using the  $\sqrt{s} = 7$  TeV and  $\sqrt{s} = 8$  TeV data set collected in 2012, within the wider framework of the phenomenological MSSM (pMSSM). The integrated luminosity was up to 20.3 fb<sup>-1</sup>. From an initial random sampling of 500 million pMSSM points, generated from the 19-parameter pMSSM, a total of 310,327 model points with  $\bar{\chi}_1^0$  LSP were selected each of which satisfies constraints from previous collider searches, precision measurements, cold dark matter energy density measurements and direct dark matter searches. The impact of the ATLAS Run 1 searches on this space was presented, considering the fraction of model points surviving, after projection into two-dimensional spaces of sparticle masses. Good complementarity is observed between different ATLAS analyses, with almost all showing regions of unique sensitivity. ATLAS searches have good sensitivity at LSP mass below 800 GeV.
- <sup>4</sup> AAD 15CA searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with one or more photons, hadronic jets or  $b$ -jets and  $E_T$ . No significant excess above the Standard Model expectations is observed. Limits are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for bino-like or higgsino-bino admixtures NLSP, see Fig. 8, 10, 11.
- <sup>5</sup> KHACHATRYAN 15AF searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least two energetic jets and significant  $E_T$ , using the transverse mass variable  $M_{T2}$  to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay  $\bar{g} \rightarrow q\bar{q}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 13(a), or where the decay  $\bar{g} \rightarrow b\bar{b}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 13(b), or where the decay  $\bar{g} \rightarrow t\bar{t}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 13(c). See also Table 5. Exclusions in the CMSSM, assuming  $\tan\beta = 30$ ,  $A_0 = -2 \max(m_0, m_{1/2})$  and  $\mu > 0$ , are also presented, see Fig. 15.
- <sup>6</sup> KHACHATRYAN 15i searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events in which  $b$ -jets and four  $W$ -bosons are produced. Five individual search channels are combined (fully hadronic, single lepton, same-sign dilepton, opposite-sign dilepton, multi-lepton). No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a simplified model where the decay  $\bar{g} \rightarrow t\bar{t}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 5. Also a simplified model with gluinos decaying into on-shell top squarks is considered, see Fig. 6.
- <sup>7</sup> KHACHATRYAN 15x searched in 19.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least two energetic jets, at least one of which is required to originate from a  $b$  quark, and significant  $E_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay  $\bar{g} \rightarrow b\bar{b}\bar{\chi}_1^0$  and the decay  $\bar{g} \rightarrow t\bar{t}\bar{\chi}_1^0$  take place with branching ratios varying between 0, 50 and 100%, see Figs. 13 and 14.
- <sup>8</sup> AAD 14AE searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for strongly produced supersymmetric particles in events containing jets and large missing transverse momentum, and no electrons or muons. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos and squarks, see Figures 5, 6 and 7. Limits are also derived in the mSUGRA/CMSSM with parameters  $\tan\beta = 30$ ,  $A_0 = -2 m_0$  and  $\mu > 0$ , see their Fig. 8.
- <sup>9</sup> AAD 14AG searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing one hadronically decaying  $\tau$ -lepton, zero or one additional light leptons (electrons or muons), jets and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set in several SUSY scenarios. For an interpretation in the minimal GMSB model, see their Fig. 8. For an interpretation in the mSUGRA/CMSSM with parameters  $\tan\beta = 30$ ,  $A_0 = -2 m_0$  and  $\mu > 0$ , see their Fig. 9. For an interpretation in the framework of natural Gauge Mediation, see Fig. 10. For an interpretation in the bRPV scenario, see their Fig. 11.
- <sup>10</sup> AAD 14x searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in an R-parity violating simplified model where the decay  $\bar{g} \rightarrow q\bar{q}\bar{\chi}_1^0$ , with  $\bar{\chi}_1^0 \rightarrow e^\pm e^\mp \nu$ , takes place with a branching ratio of 100%, see Fig. 8.
- <sup>11</sup> AAD 14x searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a general gauge-mediation model (GGM) where the decay  $\bar{g} \rightarrow q\bar{q}\bar{\chi}_1^0$ , with  $\bar{\chi}_1^0 \rightarrow e^\pm e^\mp \bar{G}$ , takes place with a branching ratio of 100%, for two choices of  $\tan\beta = 1.5$  and 30, see Fig. 11. Also some constraints on the higgsino mass parameter  $\mu$  are discussed.

- <sup>12</sup> CHATRCHYAN 14AH searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with at least two energetic jets and significant  $E_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\bar{g} \rightarrow q\bar{q}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 28. Exclusions in the CMSSM, assuming  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , are also presented, see Fig. 26.
- <sup>13</sup> CHATRCHYAN 14AH searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with at least two energetic jets and significant  $E_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a  $b$ -quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\bar{g} \rightarrow b\bar{b}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , are also presented, see Fig. 26.
- <sup>14</sup> CHATRCHYAN 14AH searched in 4.7 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with at least two energetic jets and significant  $E_T$ , using the razor variables ( $M_R$  and  $R^2$ ) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a  $b$ -quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay  $\bar{g} \rightarrow t\bar{t}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming  $\tan\beta = 10$ ,  $A_0 = 0$  and  $\mu > 0$ , are also presented, see Fig. 26.
- <sup>15</sup> CHATRCHYAN 14i searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing multijets and large  $E_T$ . No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos that decay via  $\bar{g} \rightarrow q\bar{q}\bar{\chi}_1^0$  with a 100% branching ratio, see Fig. 7b, or via  $\bar{g} \rightarrow t\bar{t}\bar{\chi}_1^0$  with a 100% branching ratio, see Fig. 7c, or via  $\bar{g} \rightarrow q\bar{q}W/Z\bar{\chi}_1^0$ , see Fig. 7d.
- <sup>16</sup> CHATRCHYAN 14N searched in 19.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing a single isolated electron or muon and multiple jets, at least two of which are identified as originating from a  $b$ -quark. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in three simplified models of gluino pair production with subsequent decay into virtual or on-shell top squarks, where each of the top squarks decays in turn into a top quark and a  $\bar{\chi}_1^0$ , see Fig. 4. The models differ in which masses are allowed to vary.
- <sup>17</sup> CHATRCHYAN 14P searched in 19.4 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for three-jet resonances produced in the decay of a gluino in R-parity violating supersymmetric models. No excess over the expected SM background is observed. Assuming a 100% branching ratio for the gluino decay into three light-flavour jets, limits are set on the cross section of gluino pair production, see Fig. 7, and gluino masses below 650 GeV are excluded at 95% C.L. Assuming a 100% branching ratio for the gluino decaying to one  $b$ -quark jet and two light-flavour jets, gluino masses between 200 GeV and 835 GeV are excluded at 95% C.L.
- <sup>18</sup> CHATRCHYAN 14R searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a slepton co-NLSP simplified model (GMSB) where the decay  $\bar{g} \rightarrow q\bar{q}e^\pm e^\mp \bar{G}$  takes place with a branching ratio of 100%, see Fig. 8.
- <sup>19</sup> CHATRCHYAN 14R searched in 19.5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a simplified model where the decay  $\bar{g} \rightarrow t\bar{t}\bar{\chi}_1^0$  takes place with a branching ratio of 100%, see Fig. 11.
- <sup>20</sup> AAD 15AB searched for the decay of neutral, weakly interacting, long-lived particles in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV. Signal events require at least two reconstructed vertices possibly originating from long-lived particles decaying to jets in the inner tracking detector and muon spectrometer. No significant excess of events over the expected background was found. Results were interpreted in Stealth SUSY benchmark models where a pair of gluinos decay to long-lived singlinos,  $\tilde{S}$ , which in turn each decay to a low-mass gravitino and a pair of jets. The 95% confidence-level limits are set on the cross section  $\times$  branching ratio for the decay  $\bar{g} \rightarrow \tilde{S}g$ , as a function of the singlino proper lifetime ( $c\tau$ ). See their Fig. 10(f).
- <sup>21</sup> AAD 15Ai searched in 20 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing at least one isolated lepton (electron or muon), jets, and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the gluino mass in the CMSSM/mSUGRA, see Fig. 15, in the NUHM2, see Fig. 16, and in various simplified models, see Figs. 18–22.
- <sup>22</sup> AAD 15CB searched for events containing at least one long-lived particle that decays at a significant distance from its production point (displaced vertex, DV) into two leptons or into five or more charged particles in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The dilepton signature is characterised by DV formed from at least two lepton candidates. Four different final states were considered for the multitrack signature, in which the DV must be accompanied by a high-transverse momentum muon or electron candidate that originates from the DV, jets or missing transverse momentum. No events were observed in any of the signal regions. Results were interpreted in SUSY scenarios involving R-parity violation, split supersymmetry, and gauge mediation. See their Fig. 12–20.
- <sup>23</sup> AAD 15x searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events containing large number of jets, no requirements on missing transverse momentum and no isolated electrons or muons. The sensitivity of the search is enhanced by considering the number of  $b$ -tagged jets and the scalar sum of masses of large-radius jets in an event. No evidence was found for excesses above the expected level of Standard Model background. Exclusion limits at 95% C.L. are set on the gluino mass assuming the gluino decays to various quark flavors, and for various neutralino masses. See their Fig. 11–16.
- <sup>24</sup> KHACHATRYAN 15AD searched in 19.4 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for events with two opposite-sign same flavor isolated leptons featuring either a kinematic edge, or a peak at the Z-boson mass, in the invariant mass spectrum. No evidence for a statistically significant excess over the expected SM backgrounds is observed and 95% C.L. exclusion limits are derived in a simplified model of gluino pair production where the gluino decays into quarks, a Z-boson, and a massless gravitino LSP, see Fig. 9.

# Searches Particle Listings

## Supersymmetric Particle Searches

- <sup>25</sup> KHACHATRYAN 15AZ searched in  $19.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with either at least one photon, hadronic jets and  $E_T$  (single photon channel) or with at least two photons and at least one jet and using the razor variables. No significant excess above the Standard Model expectations is observed. Limits are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for both a bino-like and wino-like neutralino NLSP scenario, see Fig. 8 and 9.
- <sup>26</sup> AAD 14AX searched in  $20.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for the strong production of supersymmetric particles in events containing either zero or at least one high- $p_T$  lepton, large missing transverse momentum, high jet multiplicity and at least three jets identified as originating from  $b$ -quarks. No excess over the expected SM background is observed. Limits are derived in mSUGRA/CMSSM models with  $\tan\beta = 30$ ,  $A_0 = -2m_0$  and  $\mu > 0$ , see their Fig. 14. Also, exclusion limits in simplified models containing gluinos and scalar top and bottom quarks are set, see their Figures 12, 13.
- <sup>27</sup> AAD 14E searched in  $20.3 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for strongly produced supersymmetric particles in events containing jets and two same-sign leptons or three leptons. The search also utilises jets originating from  $b$ -quarks, missing transverse momentum and other variables. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos and squarks, see Figures 5 and 6. In the  $\tilde{g} \rightarrow qq'\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^{(*)}\pm\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z^{(*)}\tilde{\chi}_1^0$  simplified model, the following assumptions have been made:  $m_{\tilde{\chi}_1^\pm} = 0.5 m_{\tilde{\chi}_1^0} + m_{\tilde{g}}, m_{\tilde{\chi}_2^0} = 0.5 (m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_1^\pm}), m_{\tilde{\chi}_1^0} < 520 \text{ GeV}$ . In the  $\tilde{g} \rightarrow qq'\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow t\bar{t}\nu\tilde{\chi}_1^0$  or  $\tilde{g} \rightarrow qq'\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow t\bar{t}\ell^+\ell^-(\nu\nu)\tilde{\chi}_1^0$  simplified model, the following assumptions have been made:  $m_{\tilde{\chi}_1^\pm} = m_{\tilde{\chi}_2^0} = 0.5 (m_{\tilde{\chi}_1^0} + m_{\tilde{g}}), m_{\tilde{\chi}_1^0} < 660 \text{ GeV}$ . Limits are also derived in the mSUGRA/CMSSM, bRPV and GMSB models, see their Fig. 8.
- <sup>28</sup> CHATRCHYAN 14H searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay  $\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, or where the decay  $\tilde{g} \rightarrow \tilde{t}\bar{t}, \tilde{t} \rightarrow t\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, with varying mass of the  $\tilde{\chi}_1^0$ , or where the decay  $\tilde{g} \rightarrow b\bar{b}, \tilde{b} \rightarrow t\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, with varying mass of the  $\tilde{\chi}_1^\pm$ , see Fig. 5.
- <sup>29</sup> CHATRCHYAN 14H searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay  $\tilde{g} \rightarrow qq'\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, with varying mass of the  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_1^0$ , see Fig. 7.
- <sup>30</sup> CHATRCHYAN 14H searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay  $\tilde{g} \rightarrow b\bar{t}\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^\pm\tilde{\chi}_1^0$  takes place with a branching ratio of 100%, for two choices of  $m_{\tilde{\chi}_1^\pm}$  and fixed  $m_{\tilde{\chi}_1^0}$ , see Fig. 6.
- <sup>31</sup> CHATRCHYAN 14H searched in  $19.5 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the R-parity violating decay  $\tilde{g} \rightarrow tbs$  takes place with a branching ratio of 100%, see Fig. 8.

### Long-lived/light $\tilde{g}$ (Gluino) MASS LIMIT

Limits on light gluinos ( $m_{\tilde{g}} < 5 \text{ GeV}$ ), or gluinos which leave the detector before decaying.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>1270	95	<sup>1</sup> AAD	15AE ATLS	$\tilde{g}$ R-hadron, generic R-hadron model
>1360	95	<sup>1</sup> AAD	15AE ATLS	$\tilde{g}$ decaying to 300 GeV stable sleptons, LeptoSUSY model
>1115	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g}$ R-hadron, stable
>1185	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow (g/q\bar{q})\tilde{\chi}_1^0$ , lifetime 10 ns, $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
>1099	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow (g/q\bar{q})\tilde{\chi}_1^0$ , lifetime 10 ns, $m_{\tilde{g}} - m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
>1182	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , lifetime 10 ns, $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
>1157	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , lifetime 10 ns, $m_{\tilde{g}} - m_{\tilde{\chi}_1^0} = 480 \text{ GeV}$
> 869	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow (g/q\bar{q})\tilde{\chi}_1^0$ , lifetime 1 ns, $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
> 821	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow (g/q\bar{q})\tilde{\chi}_1^0$ , lifetime 1 ns, $m_{\tilde{g}} - m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
> 836	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , lifetime 1 ns, $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
> 836	95	<sup>2</sup> AAD	15BMATLS	$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ , lifetime 10 ns, $m_{\tilde{g}} - m_{\tilde{\chi}_1^0} = 480 \text{ GeV}$
>1000	95	<sup>3</sup> KHACHATRY..15AK CMS		$\tilde{g}$ R-hadrons, $10 \mu\text{s} < \tau < 1000 \text{ s}$
> 880	95	<sup>3</sup> KHACHATRY..15AK CMS		$\tilde{g}$ R-hadrons, $1 \mu\text{s} < \tau < 1000 \text{ s}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 985	95	<sup>4</sup> AAD	13AA ATLS	$\tilde{g}$ , R-hadrons, generic interaction model
> 832	95	<sup>5</sup> AAD	13BC ATLS	R-hadrons, $\tilde{g} \rightarrow g/q\bar{q}\tilde{\chi}_1^0$ , generic R-hadron model, lifetime between $10^{-5}$ and $10^3 \text{ s}$ , $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
>1322	95	<sup>6</sup> CHATRCHYAN13AB CMS		long-lived $\tilde{g}$ forming R-hadrons, $f = 0.1$ , cloud interaction model
none 200–341	95	<sup>7</sup> AAD	12P ATLS	long-lived $\tilde{g} \rightarrow g\tilde{\chi}_1^0, m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$
> 640	95	<sup>8</sup> CHATRCHYAN12AN CMS		long-lived $\tilde{g} \rightarrow g\tilde{\chi}_1^0$
>1098	95	<sup>9</sup> CHATRCHYAN12L CMS		long-lived $\tilde{g}$ forming R-hadrons, $f = 0.1$
> 586	95	<sup>10</sup> AAD	11K ATLS	stable $\tilde{g}$
> 544	95	<sup>11</sup> AAD	11P ATLS	stable $\tilde{g}$ , GMSB scenario, $\tan\beta=5$
> 370	95	<sup>12</sup> KHACHATRY..11 CMS		long lived $\tilde{g}$
> 398	95	<sup>13</sup> KHACHATRY..11C CMS		stable $\tilde{g}$

<sup>1</sup> AAD 15AE searched in  $19.1 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ATLAS muon system. In the absence of an excess of events above the expected backgrounds, limits are set R-hadrons in various scenarios, see Fig. 11. Limits are also set in LeptoSUSY models where the gluino decays to stable 300 GeV leptons, see Fig. 9.

<sup>2</sup> AAD 15BM searched in  $18.4 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for stable and metastable non-relativistic charged particles through their anomalous specific ionization energy loss in the ATLAS pixel detector. In absence of an excess of events above the expected backgrounds, limits are set within a generic R-hadron model, on stable gluino R-hadrons (see Table 5) and on metastable gluino R-hadrons decaying to  $(g/q\bar{q})$  plus a light  $\tilde{\chi}_1^0$  (see Fig. 7) and decaying to  $t\bar{t}$  plus a light  $\tilde{\chi}_1^0$  (see Fig. 9).

<sup>3</sup> KHACHATRYAN 15AK looked in a data set corresponding to  $18.6 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$ , and a search interval corresponding to 281 h of trigger lifetime, for long-lived particles that have stopped in the CMS detector. No evidence for an excess over the expected background in a cloud interaction model is observed. Assuming the decay  $\tilde{g} \rightarrow g\tilde{\chi}_1^0$  and lifetimes between 1  $\mu\text{s}$  and 1000 s, limits are derived on  $\tilde{g}$  production as a function of  $m_{\tilde{\chi}_1^0}$ , see Figs. 4 and 6. The exclusions require that  $m_{\tilde{\chi}_1^0}$  is kinematically consistent with the minimum values of the jet energy thresholds used.

<sup>4</sup> AAD 13AA searched in  $4.7 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events containing colored long-lived particles that hadronize forming R-hadrons. No significant excess above the expected background was found. Long-lived R-hadrons containing a  $\tilde{g}$  are excluded for masses up to 985 GeV at 95% C.L. in a general interaction model. Also, limits independent of the fraction of R-hadrons that arrive charged in the muon system were derived, see Fig. 6.

<sup>5</sup> AAD 13BC searched in  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  and in  $22.9 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for bottom squark R-hadrons that have come to rest within the ATLAS calorimeter and decay at some later time to hadronic jets and a neutralino. In absence of an excess of events above the expected backgrounds, limits are set on gluino masses for different decays, lifetimes, and neutralino masses, see their Table 6 and Fig. 10.

<sup>6</sup> CHATRCHYAN 13AB looked in  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  and in  $18.8 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of  $\tilde{g}$ 's. No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 8 and Table 5), depending on the fraction,  $f$ , of formation of  $\tilde{g}$ - $\tilde{g}$  (R-gluonball) states. The quoted limit is for  $f = 0.1$ , while for  $f = 0.5$  it degrades to 1276 GeV. In the conservative scenario where every hadronic interaction causes it to become neutral, the limit decreases to 928 GeV for  $f = 0.1$ .

<sup>7</sup> AAD 12P looked in  $31 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with pair production of long-lived gluinos. The hadronization of the gluinos leads to R-hadrons which may stop inside the detector and later decay via  $\tilde{g} \rightarrow g\tilde{\chi}_1^0$  during gaps between the proton bunches. No significant excess over the expected background is observed. From a counting experiment, a limit at 95% C.L. on the cross section as a function of  $m_{\tilde{g}}$  is derived for  $m_{\tilde{\chi}_1^0} = 100 \text{ GeV}$ , see Fig. 4. The limit is valid for lifetimes between  $10^{-5}$  and  $10^3$  seconds and assumes the Generic matter interaction model for the production cross section.

<sup>8</sup> CHATRCHYAN 12AN looked in  $4.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with pair production of long-lived gluinos. The hadronization of the gluinos leads to R-hadrons which may stop inside the detector and later decay via  $\tilde{g} \rightarrow g\tilde{\chi}_1^0$  during gaps between the proton bunches. No significant excess over the expected background is observed. From a counting experiment, a limit at 95% C.L. on the cross section as a function of  $m_{\tilde{g}}$  is derived, see Fig. 3. The mass limit is valid for lifetimes between  $10^{-5}$  and  $10^3$  seconds, for what they call "the daughter gluon energy  $E_g > 100 \text{ GeV}$  and assuming the cloud interaction model for R-hadrons. Supersedes KHACHATRYAN 11.

<sup>9</sup> CHATRCHYAN 12L looked in  $5.0 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of  $\tilde{g}$ 's. No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 3), depending on the fraction,  $f$ , of formation of  $\tilde{g}$ - $\tilde{g}$  (R-gluonball) states. The quoted limit is for  $f = 0.1$ , while for  $f = 0.5$  it degrades to 1046 GeV. In the conservative scenario where every hadronic interaction causes it to become neutral, the limit decreases to 928 GeV for  $f = 0.1$ . Supersedes KHACHATRYAN 11C.

<sup>10</sup> AAD 11K looked in  $34 \text{ pb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or time of flight in the tile calorimeter, from pair production of  $\tilde{g}$ . No evidence for an excess over the SM expectation is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 4), for a fraction,  $f = 10\%$ , of formation of  $\tilde{g}$ - $\tilde{g}$  (R-gluonball). If instead of a phase space driven approach for the hadronic scattering of the R-hadrons,

a triple-Regge model or a bag-model is used, the limit degrades to 566 and 562 GeV, respectively.

- <sup>11</sup> AAD 11P looked in 37 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with heavy stable particles, reconstructed and identified by their time of flight in the Muon System. There is no requirement on their observation in the tracker to increase the sensitivity to cases where gluinos have a large fraction,  $f$ , of formation of neutral  $\tilde{g} - g$  (R-gluonball). No evidence for an excess over the SM expectation is observed. Limits are derived as a function of mass (see Fig. 4), for  $f=0.1$ . For fractions  $f = 0.5$  and 1.0 the limit degrades to 537 and 530 GeV, respectively.

- <sup>12</sup> KHACHATRYAN 11 looked in 10 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with pair production of long-lived gluinos. The hadronization of the gluinos leads to R-hadrons which may stop inside the detector and later decay via  $\tilde{g} \rightarrow g \chi_1^0$  during gaps between the proton bunches. No significant excess over the expected background is observed. From a counting experiment, a limit at 95% C.L. on the cross section times branching ratio is derived for  $m_{\tilde{g}} - m_{\chi_1^0} > 100$  GeV, see their Fig. 2. Assuming 100% branching

ratio, lifetimes between 75 ns and  $3 \times 10^5$  s are excluded for  $m_{\tilde{g}} = 300$  GeV. The  $\tilde{g}$  mass exclusion is obtained with the same assumptions for lifetimes between 10  $\mu$ s and 1000 s, but shows some dependence on the model for R-hadron interactions with matter, illustrated in Fig. 3. From a time-profile analysis, the mass exclusion is 382 GeV for a lifetime of 10  $\mu$ s under the same assumptions as above.

- <sup>13</sup> KHACHATRYAN 11C looked in 3.1 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of  $\tilde{g}$ . No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 3), depending on the fraction,  $f$ , of formation of  $\tilde{g} - g$  (R-gluonball). The quoted limit is for  $f=0.1$ , while for  $f=0.5$  it degrades to 357 GeV. In the conservative scenario where every hadronic interaction causes it to become neutral, the limit decreases to 311 GeV for  $f=0.1$ .

### LIGHT $\tilde{G}$ (Gravitino) MASS LIMITS FROM COLLIDER EXPERIMENTS

The following are bounds on light ( $< 1$  eV) gravitino indirectly inferred from its coupling to matter suppressed by the gravitino decay constant.

Unless otherwise stated, all limits assume that other supersymmetric particles besides the gravitino are too heavy to be produced. The gravitino is assumed to be undetected and to give rise to a missing energy ( $\cancel{E}$ ) signature.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$> 3.5 \times 10^{-4}$	95	<sup>1</sup> AAD	15BH ATLS	jet + $\cancel{E}_T$ , $pp \rightarrow (\tilde{q}/\tilde{g}) \tilde{G}$ , $m_{\tilde{q}} = m_{\tilde{g}} = 500$ GeV
$> 3 \times 10^{-4}$	95	<sup>1</sup> AAD	15BH ATLS	jet + $\cancel{E}_T$ , $pp \rightarrow (\tilde{q}/\tilde{g}) \tilde{G}$ , $m_{\tilde{q}} = m_{\tilde{g}} = 1000$ GeV
$> 2 \times 10^{-4}$	95	<sup>1</sup> AAD	15BH ATLS	jet + $\cancel{E}_T$ , $pp \rightarrow (\tilde{q}/\tilde{g}) \tilde{G}$ , $m_{\tilde{q}} = m_{\tilde{g}} = 1500$ GeV
$> 1.09 \times 10^{-5}$	95	<sup>2</sup> ABDALLAH	05B DLPH	$e^+e^- \rightarrow \tilde{G} \tilde{G} \gamma$
$> 1.35 \times 10^{-5}$	95	<sup>3</sup> ACHARD	04E L3	$e^+e^- \rightarrow \tilde{G} \tilde{G} \gamma$
$> 1.3 \times 10^{-5}$		<sup>4</sup> HEISTER	03C ALEP	$e^+e^- \rightarrow \tilde{G} \tilde{G} \gamma$
$> 11.7 \times 10^{-6}$	95	<sup>5</sup> ACOSTA	02H CDF	$p\bar{p} \rightarrow \tilde{G} \tilde{G} \gamma$
$> 8.7 \times 10^{-6}$	95	<sup>6</sup> ABBIENDI,G	00D OPAL	$e^+e^- \rightarrow \tilde{G} \tilde{G} \gamma$

- <sup>1</sup> AAD 15BH searched in 20.3 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 8$  TeV for associated production of a light gravitino and a squark or gluino. The squark (gluino) is assumed to decay exclusively to a quark (gluon) and a gravitino. No evidence was found for an excess above the expected level of Standard Model background and 95% C.L. lower limits were set on the gravitino mass as a function of the squark/gluino mass, both in the case of degenerate and non-degenerate squark/gluino masses, see Figs. 14 and 15.

- <sup>2</sup> ABDALLAH 05B use data from  $\sqrt{s} = 180$ –208 GeV. They look for events with a single photon +  $\cancel{E}$  final states from which a cross section limit of  $\sigma < 0.18$  pb at 208 GeV is obtained, allowing a limit on the mass to be set. Supersedes the results of ABREU 00Z.

- <sup>3</sup> ACHARD 04E use data from  $\sqrt{s} = 189$ –209 GeV. They look for events with a single photon +  $\cancel{E}$  final states from which a limit on the Gravitino mass is set corresponding to  $\sqrt{\mathcal{F}} > 238$  GeV. Supersedes the results of ACCIARRI 99R.

- <sup>4</sup> HEISTER 03C use the data from  $\sqrt{s} = 189$ –209 GeV to search for  $\gamma \cancel{E}_T$  final states.

- <sup>5</sup> ACOSTA 02H looked in 87 pb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s}=1.8$  TeV for events with a high- $E_T$  photon and  $\cancel{E}_T$ . They compared the data with a GMSB model where the final state could arise from  $q\bar{q} \rightarrow \tilde{G} \tilde{G} \gamma$ . Since the cross section for this process scales as  $1/|F|^4$ , a limit at 95% CL is derived on  $|F|^{1/2} > 221$  GeV. A model independent limit for the above topology is also given in the paper.

- <sup>6</sup> ABBIENDI,G 00D searches for  $\gamma \cancel{E}$  final states from  $\sqrt{s}=189$  GeV.

### Supersymmetry Miscellaneous Results

Results that do not appear under other headings or that make nonminimal assumptions.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C **38** 070001 (2014) (<http://pdg.lbl.gov>).

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>1</sup> AAD	13P ATLS	dark $\gamma$ , hidden valley
		<sup>2</sup> AALTONEN	12AB CDF	hidden-valley Higgs
none 100–185	95	<sup>3</sup> AAD	11AA ATLS	scalar gluons
		<sup>4</sup> CHATRCHYAN 11E	CMS	$\mu\mu$ resonances
		<sup>5</sup> ABAZOV	10N D0	$\gamma_D$ , hidden valley

- <sup>1</sup> AAD 13P searched in 5 fb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for single lepton-jets with at least four muons; pairs of lepton-jets, each with two or more muons; and pairs of lepton-jets with two or more electrons. All of these could be signatures of Hidden Valley supersymmetric models. No statistically significant deviations from the Standard Model expectations are found. 95% C.L. limits are placed on the production cross section times branching ratio of dark photons for several parameter sets of a Hidden Valley model.

- <sup>2</sup> AALTONEN 12AB looked in 5.1 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for anomalous production of multiple low-energy leptons in association with a  $W$  or  $Z$  boson. Such events may occur in hidden valley models in which a supersymmetric Higgs boson is produced in association with a  $W$  or  $Z$  boson, with  $H \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0$  pair and with the  $\tilde{\chi}_1^0$  further decaying into a dark photon ( $\gamma_D$ ) and the unobservable lightest SUSY particle of the hidden sector. As the  $\gamma_D$  is expected to be light, it may decay into a lepton pair. No significant excess over the SM expectation is observed and a limit at 95% C.L. is set on the cross section for a benchmark model of supersymmetric hidden-valley Higgs production.

- <sup>3</sup> AAD 11AA looked in 34 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with  $\geq 4$  jets originating from pair production of scalar gluons, each decaying to two gluons. No two-jet resonances are observed over the SM background. Limits are derived on the cross section times branching ratio (see Fig. 3). Assuming 100% branching ratio for the decay to two gluons, the quoted exclusion range is obtained, except for a 5 GeV mass window around 140 GeV.

- <sup>4</sup> CHATRCHYAN 11E looked in 35 pb<sup>-1</sup> of  $pp$  collisions at  $\sqrt{s} = 7$  TeV for events with collimated  $\mu$  pairs (leptonic jets) from the decay of hidden sector states. No evidence for new resonance production is found. Limits are derived and compared to various SUSY models (see Fig. 4) where the LSP, either the  $\tilde{\chi}_1^0$  or a  $\tilde{q}$ , decays to dark sector particles.

- <sup>5</sup> ABAZOV 10N looked in 5.8 fb<sup>-1</sup> of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV for events from hidden valley models in which a  $\tilde{\chi}_1^0$  decays into a dark photon,  $\gamma_D$ , and the unobservable lightest SUSY particle of the hidden sector. As the  $\gamma_D$  is expected to be light, it may decay into a tightly collimated lepton pair, called lepton jet. They searched for events with  $\cancel{E}_T$  and two isolated lepton jets observable by an opposite charged lepton pair  $e e$ ,  $e \mu$  or  $\mu \mu$ . No significant excess over the SM expectation is observed, and a limit at 95% C.L. on the cross section times branching ratio is derived, see their Table I. They also examined the invariant mass of the lepton jets for a narrow resonance, see their Fig. 4, but found no evidence for a signal.

### REFERENCES FOR Supersymmetric Particle Searches

AAD	15AB	PR D92 012010	G. Aad et al.	(ATLAS Collab.)
AAD	15AE	JHEP 1501 068	G. Aad et al.	(ATLAS Collab.)
AAD	15AI	JHEP 1504 116	G. Aad et al.	(ATLAS Collab.)
AAD	15BA	EPJ C75 208	G. Aad et al.	(ATLAS Collab.)
AAD	15BG	EPJ C75 318	G. Aad et al.	(ATLAS Collab.)
Also	EPJ C75 463	G. Aad et al.	(ATLAS Collab.)	
AAD	15BH	EPJ C75 299	G. Aad et al.	(ATLAS Collab.)
AAD	15BM	EPJ C75 407	G. Aad et al.	(ATLAS Collab.)
AAD	15BV	JHEP 1510 054	G. Aad et al.	(ATLAS Collab.)
AAD	15BX	JHEP 1510 134	G. Aad et al.	(ATLAS Collab.)
AAD	15CA	PR D92 072001	G. Aad et al.	(ATLAS Collab.)
AAD	15CB	PR D92 072004	G. Aad et al.	(ATLAS Collab.)
AAD	15CJ	EPJ C75 510	G. Aad et al.	(ATLAS Collab.)
AAD	15CS	PR D91 012008	G. Aad et al.	(ATLAS Collab.)
Also	PR D92 059903 (err.)	G. Aad et al.	(ATLAS Collab.)	
AAD	15J	PRL 114 142001	G. Aad et al.	(ATLAS Collab.)
AAD	15K	PRL 114 161801	G. Aad et al.	(ATLAS Collab.)
AAD	15O	PRL 115 031801	G. Aad et al.	(ATLAS Collab.)
AAD	15X	PR D91 112016	G. Aad et al.	(ATLAS Collab.)
AJAJ	15BD	EPJ C75 595	R. Aaij et al.	(LHCb Collab.)
AARTSEN	15C	EPJ C75 20	M. G. Aartsen et al.	(IceCube Collab.)
AARTSEN	15E	EPJ C75 492	M. G. Aartsen et al.	(IceCube Collab.)
ABRAMOWSKI	15	PRL 114 081301	A. Abramowski et al.	(H.E.S.S. Collab.)
ACKERMANN	15	PR D91 122002	M. Ackermann et al.	(Fermi-LAT Collab.)
ACKERMANN	15A	JCAP 1509 008	M. Ackermann et al.	(Fermi-LAT Collab.)
ACKERMANN	15B	PRL 115 231301	M. Ackermann et al.	(Fermi-LAT Collab.)
ADRIAN-MAR.	15	JCAP 1510 068	S. Adrian-Marinez et al.	(ANTARES Collab.)
AGNES	15	PL B743 456	P. Agnes et al.	(DarkSide-50 Collab.)
AGNESE	15A	PR D91 052021	R. Agnese et al.	(SuperCDMS Collab.)
AGNESE	15B	PR D92 072003	R. Agnese et al.	(SuperCDMS Collab.)
AMOLE	15	PRL 114 231302	C. Amole et al.	(PICO Collab.)
BUCKLEY	15	PR D91 102001	M. R. Buckley et al.	(Super-Kamiokande Collab.)
CHOI	15	PRL 114 141301	K. Choi et al.	(CMS Collab.)
KHACHATRYAN	15AB	JHEP 1501 096	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AD	JHEP 1504 124	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AF	JHEP 1505 078	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AH	JHEP 1506 116	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AK	EPJ C75 151	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AO	EPJ C75 325	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AR	PL B743 503	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AZ	PR D92 072006	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15E	PRL 114 061801	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15I	PL B745 5	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15L	PL B747 38	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15O	PL B748 255	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15W	PR D91 052012	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15X	PR D91 052018	V. Khachatryan et al.	(CMS Collab.)
ROLBIECKI	15	PL B750 247	K. Rolbiecki, J. Tattersall	(MADE, HEID)
XIAO	15	PR D92 052004	X. Xiao et al.	(PandaX Collab.)
AAD	14AE	JHEP 1409 176	G. Aad et al.	(ATLAS Collab.)
AAD	14AG	JHEP 1409 103	G. Aad et al.	(ATLAS Collab.)
AAD	14AJ	JHEP 1409 015	G. Aad et al.	(ATLAS Collab.)
AAD	14AV	JHEP 1410 096	G. Aad et al.	(ATLAS Collab.)
AAD	14AX	JHEP 1410 024	G. Aad et al.	(ATLAS Collab.)
AAD	14B	EPJ C74 2863	G. Aad et al.	(ATLAS Collab.)
AAD	14BD	JHEP 1411 118	G. Aad et al.	(ATLAS Collab.)
AAD	14BH	PR D90 112005	G. Aad et al.	(ATLAS Collab.)
AAD	14E	JHEP 1406 035	G. Aad et al.	(ATLAS Collab.)
AAD	14F	JHEP 1406 124	G. Aad et al.	(ATLAS Collab.)
AAD	14G	JHEP 1405 071	G. Aad et al.	(ATLAS Collab.)
AAD	14H	JHEP 1404 169	G. Aad et al.	(ATLAS Collab.)
AAD	14K	PR D90 012004	G. Aad et al.	(ATLAS Collab.)
AAD	14T	PR D90 052008	G. Aad et al.	(ATLAS Collab.)
AAD	14X	PR D90 052001	G. Aad et al.	(ATLAS Collab.)
AALTONEN	14	PR D90 012011	T. Aaltonen et al.	(CDF Collab.)
ACKERMANN	14	PR D89 042001	M. Ackermann et al.	(Fermi-LAT Collab.)
AGNESE	14	PRL 112 241302	R. Agnese et al.	(SuperCDMS Collab.)
AKERIB	14	PRL 112 091303	D. S. Akerib et al.	(LUX Collab.)
ALEKSIK	14	JCAP 1402 008	J. Aleksić et al.	(MAGIC Collab.)
AVRORIN	14	ASP 62 12	A. D. Avrorin et al.	(BAIKAL Collab.)
BUCHMUELLER	14	EPJ C74 2809	O. Buchmüller et al.	(CMS Collab.)
BUCHMUELLER	14A	EPJ C74 2922	O. Buchmüller et al.	(CMS Collab.)
CHATRCHYAN	14AH	PR D90 112001	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	14H	JHEP 1401 163	S. Chatrchyan et al.	(CMS Collab.)

# Searches Particle Listings

## Supersymmetric Particle Searches

CHATRCHYAN	141	JHEP 1406 055	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	CHATRCHYAN	11B	JHEP 1106 093	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	14N	PL B733 328	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	CHATRCHYAN	11D	JHEP 1107 113	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	14P	PL B730 193	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	CHATRCHYAN	11E	JHEP 1107 098	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	14R	PR D90 032006	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	CHATRCHYAN	11V	PL B704 411	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	14U	PRL 112 161802	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	KHACHATRY...	11	PRL 106 011801	V. Khachatryan <i>et al.</i>	(CMS Collab.)
CZAKON	14	PRL 113 201803	M. Czakon <i>et al.</i>	(AACH, CAMB, UCB, LBL+)	KHACHATRY...	11C	JHEP 1103 024	V. Khachatryan <i>et al.</i>	(CMS Collab.)
FELIZARDO	14	PR D89 072013	M. Felizardo <i>et al.</i>	(SIMPLE Collab.)	ROSZKOWSKI	11	PR D83 015014	L. Roszkowski <i>et al.</i>	(CDF Collab.)
KHACHATRY...	14C	PL B736 375	V. Khachatryan <i>et al.</i>	(CMS Collab.)	AALTONEN	10	PRL 104 011801	T. Aaltonen <i>et al.</i>	(CDF Collab.)
KHACHATRY...	14I	EPJ C74 3036	V. Khachatryan <i>et al.</i>	(CMS Collab.)	AALTONEN	10R	PRL 105 081802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
KHACHATRY...	14L	PR D90 092007	V. Khachatryan <i>et al.</i>	(CMS Collab.)	AALTONEN	10Z	PRL 105 191801	T. Aaltonen <i>et al.</i>	(CDF Collab.)
KHACHATRY...	14T	PL B739 229	V. Khachatryan <i>et al.</i>	(CMS Collab.)	ABAZOV	10L	PL B693 95	V.M. Abazov <i>et al.</i>	(DO Collab.)
PDG	14	CPC 38 070001	K. Olive <i>et al.</i>	(PDG Collab.)	ABAZOV	10M	PRL 105 191802	V.M. Abazov <i>et al.</i>	(DO Collab.)
ROSZKOWSKI	14	JHEP 1408 067	L. Roszkowski, E.M. Sessolo, A.J. Williams	(WINR)	ABAZOV	10N	PRL 105 211802	V.M. Abazov <i>et al.</i>	(DO Collab.)
AAD	13	PL B718 841	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABAZOV	10P	PRL 105 221802	V.M. Abazov <i>et al.</i>	(DO Collab.)
AAD	13AA	PL B720 277	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABDO	10	JCAP 1004 014	A.A. Abdo <i>et al.</i>	(Fermi-LAT Collab.)
AAD	13AI	PL B723 15	G. Aad <i>et al.</i>	(ATLAS Collab.)	ACKERMANN	10	JCAP 1005 025	M. Ackermann	(Fermi-LAT Collab.)
AAD	13AP	PR D88 012001	G. Aad <i>et al.</i>	(ATLAS Collab.)	AHMED	10	SCI 327 1619	Z. Ahmed <i>et al.</i>	(CDMS II Collab.)
AAD	13AU	JHEP 1310 189	G. Aad <i>et al.</i>	(ATLAS Collab.)	ARMENGAUD	10	PL B687 294	E. Armengaud <i>et al.</i>	(EDELWEISS II Collab.)
AAD	13B	PL B718 879	G. Aad <i>et al.</i>	(ATLAS Collab.)	ELLIS	10	EPJ C69 201	J. Ellis, A. Mustafayev, K. Olive	(DO Collab.)
AAD	13BC	PR D88 112003	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABAZOV	09M	PRL 102 161802	V.M. Abazov <i>et al.</i>	(DO Collab.)
AAD	13BD	PR D88 112006	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABBASI	09B	PRL 102 201302	R. Abbasi <i>et al.</i>	(IceCube Collab.)
AAD	13H	JHEP 1301 131	G. Aad <i>et al.</i>	(ATLAS Collab.)	AHMED	09	PRL 102 011301	Z. Ahmed <i>et al.</i>	(CDMS Collab.)
AAD	13L	PR D87 012008	G. Aad <i>et al.</i>	(ATLAS Collab.)	ANGLOHER	09	ASP 31 270	G. Angloher <i>et al.</i>	(CREST Collab.)
AAD	13P	PL B719 299	G. Aad <i>et al.</i>	(ATLAS Collab.)	ARCHAMBAU...	09	PL B682 185	S. Archambault <i>et al.</i>	(PICASSO Collab.)
AAD	13Q	PL B719 261	G. Aad <i>et al.</i>	(ATLAS Collab.)	BUCHMUELL...	09	EPJ C64 391	O. Buchmueller <i>et al.</i>	(LOIC, FNAL, CERN+)
AAD	13R	PL B719 280	G. Aad <i>et al.</i>	(ATLAS Collab.)	DREINER	09	EPJ C62 547	H. Dreiner <i>et al.</i>	
AALSETH	13	PR D88 012002	C.E. Aalseth <i>et al.</i>	(CoGENT Collab.)	LEBEDENKO	09	PR D80 052010	V.N. Lebedenko <i>et al.</i>	(ZEPLIN-III Collab.)
AALTONEN	13I	PR D88 031103	T. Aaltonen <i>et al.</i>	(CDF Collab.)	LEBEDENKO	09A	PRL 103 151302	V.N. Lebedenko <i>et al.</i>	(ZEPLIN-III Collab.)
AALTONEN	13Q	PRL 110 201802	T. Aaltonen <i>et al.</i>	(CDF Collab.)	SORENSEN	09	NIM A601 339	P. Sorensen <i>et al.</i>	(XENON10 Collab.)
AARTSEN	13	PRL 110 131302	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)	ABAZOV	08F	PL B659 856	V.M. Abazov <i>et al.</i>	(DO Collab.)
AARTSEN	13C	PR D88 122001	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)	ABAZOV	07M	PRL 100 021303	J. Angle <i>et al.</i>	(XENON10 Collab.)
ABAZOV	13B	PR D87 052011	V.M. Abazov <i>et al.</i>	(DO Collab.)	ANGLE	08A	PRL 101 091301	J. Angle <i>et al.</i>	(XENON10 Collab.)
ABRAMOWSKI	13	PRL 110 041301	A. Abramowski <i>et al.</i>	(H.E.S.S. Collab.)	BEDNYAKOV	08	PAN 71 111	V.A. Bednyakov, H.P. Klapdor-Kleingrothaus, I.V. Krivosheina	
ACKERMANN	13A	PR D88 082002	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)	BEHNKE	08	SCI 319 933	E. Behnke	(COUPP Collab.)
ADRIAN-MAR...	13	JCAP 1311 032	R. Adrian-Martinez <i>et al.</i>	(ANTARES Collab.)	BENETTI	08	ASP 28 495	P. Benetti <i>et al.</i>	(WARP Collab.)
AGNESE	13	PR D88 031104	R. Agnese <i>et al.</i>	(CDMS Collab.)	BUCHMUELL...	08	JHEP 0809 117	O. Buchmueller <i>et al.</i>	
AGNESE	13A	PRL 111 251301	R. Agnese <i>et al.</i>	(CDMS Collab.)	ELLIS	08	PR D78 075012	J. Ellis, K. Olive, P. Sandick	(CERN, MINN)
APRILE	13	PRL 111 021301	E. Aprile <i>et al.</i>	(XENON100 Collab.)	ABULENCIA	07H	PRL 98 131804	A. Abulencia <i>et al.</i>	(CDF Collab.)
BERGSTROM	13	PRL 111 171101	L. Bergstrom <i>et al.</i>		ALNER	07A	ASP 28 287	G.J. Alner <i>et al.</i>	(ZEPLIN-II Collab.)
BOLIEV	13	JCAP 1309 019	M. Boliev <i>et al.</i>		CALIBBI	07	JHEP 0709 081	L. Calibbi <i>et al.</i>	
CABRERA	13	JHEP 1307 182	M. Cabrera, J. Casas, R. de Austri	(CMS Collab.)	ELLIS	07	JHEP 0706 079	J. Ellis, K. Olive, P. Sandick	(CERN, MINN)
CHATRCHYAN	13	PL B718 815	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	LEE	07A	PRL 99 091301	H.S. Lee <i>et al.</i>	(KIMS Collab.)
CHATRCHYAN	13AB	JHEP 1307 122	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ABBIENDI	06B	EPJ C46 307	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
CHATRCHYAN	13AH	PL B722 273	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ACHTERBERG	06	ASP 26 129	A. Achterberg <i>et al.</i>	(AMANDA Collab.)
CHATRCHYAN	13AO	PR D87 072001	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ACKERMANN	06	ASP 24 459	M. Ackermann <i>et al.</i>	(AMANDA Collab.)
CHATRCHYAN	13AT	PR D88 052017	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	AKERIB	06	PR D73 011102	D.S. Akerib <i>et al.</i>	(CDMS Collab.)
CHATRCHYAN	13AV	PRL 111 081802	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	AKERIB	06A	PRL 96 011302	D.S. Akerib <i>et al.</i>	(CDMS Collab.)
CHATRCHYAN	13G	JHEP 1301 077	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ALLANACH	06	PR D73 015013	B.C. Allanach <i>et al.</i>	
CHATRCHYAN	13H	PL B719 42	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	BEHOIT	06	PL B637 156	A. Benoit <i>et al.</i>	
CHATRCHYAN	13T	EPJ C73 2568	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	DE-AUSTRI	06	JHEP 0605 002	R.R. de Austri, R. Trotta, L. Roszkowski	
CHATRCHYAN	13V	JHEP 1303 037	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	DEBOER	06	PL B636 13	W. de Boer <i>et al.</i>	
Also		JHEP 1307 041 (err.)	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	LEP-SLC	06	PRPL 427 257	ALEPH, DELPHI, L3, OPAL, SLD and working groups	
CHATRCHYAN	13W	JHEP 1303 111	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	SHIMIZU	06A	PL B633 195	Y. Shimizu <i>et al.</i>	
ELLIS	13B	EPJ C73 2403	J. Ellis <i>et al.</i>		SMITH	06	PL B642 567	N.J.T. Smith, A.S. Murphy, T.J. Summer	
JIN	13	JCAP 1311 026	H.-B. Jin, Y.-L. Wu, Y.-F. Zhou		ABAZOV	05A	PRL 94 041801	V.M. Abazov <i>et al.</i>	(DO Collab.)
KOPP	13	PR D88 076013	J. Kopp		ABDALLAH	05B	EPJ C38 395	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
LI	13B	PRL 110 261301	H.B. Li <i>et al.</i>	(TEXONO Collab.)	AKERIB	05	PR D72 052009	D.S. Akerib <i>et al.</i>	(CDMS Collab.)
STREGE	13	JCAP 1304 013	C. Strege <i>et al.</i>		ALNER	05	PL B616 17	G.J. Alner <i>et al.</i>	(UK Dark Matter Collab.)
AAD	12AF	PL B714 180	G. Aad <i>et al.</i>	(ATLAS Collab.)	ALNER	05A	ASP 23 444	G.J. Alner <i>et al.</i>	(UK Dark Matter Collab.)
AAD	12AG	PL B714 197	G. Aad <i>et al.</i>	(ATLAS Collab.)	ANGLOHER	05	ASP 23 325	G. Angloher <i>et al.</i>	(CREST-II Collab.)
AAD	12AN	PRL 108 181802	G. Aad <i>et al.</i>	(ATLAS Collab.)	BAER	05	JHEP 0507 065	H. Baer <i>et al.</i>	(FSU, MSU, HAWA)
AAD	12AS	PRL 108 261804	G. Aad <i>et al.</i>	(ATLAS Collab.)	BARNABE-HE...	05	PL B624 186	M. Barnabe-Heider <i>et al.</i>	(PICASSO Collab.)
AAD	12AX	PR D85 012006	G. Aad <i>et al.</i>	(ATLAS Collab.)	ELLIS	05	PR D71 095007	J. Ellis <i>et al.</i>	
Also		PR D87 099003 (err.)	G. Aad <i>et al.</i>	(ATLAS Collab.)	SANGLARD	05	PR D71 122002	V. Sanglard <i>et al.</i>	(EDELWEISS Collab.)
AAD	12BJ	EPJ C72 1933	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABBIENDI	04	EPJ C32 453	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
AAD	12CJ	PR D86 092002	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABBIENDI	04F	EPJ C33 149	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
AAD	12CM	EPJ C72 2215	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABBIENDI	04H	EPJ C35 1	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
AAD	12CP	PL B718 411	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABBIENDI	04N	PL B602 167	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
AAD	12CT	JHEP 1212 124	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABDALLAH	04H	EPJ C34 145	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
AAD	12P	EPJ C72 1965	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABDALLAH	04M	EPJ C36 1	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
AAD	12R	PL B707 478	G. Aad <i>et al.</i>	(ATLAS Collab.)	Also		EPJ C37 129 (err.)	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
AAD	12T	PL B709 137	G. Aad <i>et al.</i>	(ATLAS Collab.)	ACHARD	04	PL B580 37	P. Achard <i>et al.</i>	(L3 Collab.)
AAD	12W	PL B710 67	G. Aad <i>et al.</i>	(ATLAS Collab.)	ACHARD	04E	PL B587 16	P. Achard <i>et al.</i>	(L3 Collab.)
AALTONEN	12AB	PR D85 092001	T. Aaltonen <i>et al.</i>	(CDF Collab.)	AKERIB	04	PRL 93 211301	D. Akerib <i>et al.</i>	(CDMSII Collab.)
ABAZOV	12AD	PR D86 071701	V.M. Abazov <i>et al.</i>	(CDF Collab.)	BALTZ	04	JHEP 0410 052	E. Baltz, P. Gondolo	
ABBASI	12	PR D85 042002	R. Abbasi <i>et al.</i>	(IceCube Collab.)	BELANGER	04	JHEP 0403 012	G. Belanger <i>et al.</i>	
AKIMOV	12	PL B709 14	D.Yu. Akimov <i>et al.</i>	(ZEPLIN-III Collab.)	BOTTINO	04	PR D69 037302	A. Bottino <i>et al.</i>	
AKULA	12	PR D85 075001	S. Akula <i>et al.</i>	(NEAS, MICH)	DESAI	04	PR D70 083523	S. Desai <i>et al.</i>	(Super-Kamiokande Collab.)
ANGLOHER	12	EPJ C72 1971	G. Angloher <i>et al.</i>	(CREST-II Collab.)	ELLIS	04	PR D69 015005	J. Ellis <i>et al.</i>	
APRILE	12	PRL 109 181301	E. Aprile <i>et al.</i>	(XENON100 Collab.)	ELLIS	04B	PR D70 055005	J. Ellis <i>et al.</i>	
ARBEY	12A	PL B708 162	A. Arbey <i>et al.</i>		HEISTER	04	PL B583 247	A. Heister <i>et al.</i>	(ALEPH Collab.)
ARCHAMBAU...	12	PL B711 153	S. Archambault <i>et al.</i>	(PICASSO Collab.)	PIERCE	04D	PR D70 075006	A. Pierce	
BAER	12	JHEP 1205 091	H. Baer, V. Barger, A. Mustafayev	(OKLA, WISC+)	ABBIENDI	03L	PL B572 8	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
BALAZS	12	EPJ C73 2563	C. Balazs <i>et al.</i>		ABDALLAH	03D	EPJ C27 153	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
BECHTLE	12	JHEP 1206 098	P. Bechtle <i>et al.</i>	(COUPP Collab.)	ABDALLAH	03M	EPJ C31 421	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
BEHNKE	12	PR D86 052001	E. Behnke <i>et al.</i>	(COUPP Collab.)	AHMED	03	ASP 19 691	B. Ahmed <i>et al.</i>	(UK Dark Matter Collab.)
Also		PR D90 079902 (err.)	C. Beskidt <i>et al.</i>	(KARLE, JINR, ITEP)	AKERIB	03	PR D68 082002	D. Akerib <i>et al.</i>	(CDMS Collab.)
BESKIDT	12	EPJ C72 2166	C. Beskidt <i>et al.</i>	(TORI, SOGA)	BAER	03A	JCAP 0309 007	H. Baer <i>et al.</i>	
BOTTINO	12	PR D85 095013	A. Bottino, N. Fornengo, S. Scopel		BOTTINO	03	PR D68 043506	A. Bottino <i>et al.</i>	
BUCHMUELL...	12	EPJ C72 2020	O. Buchmueller <i>et al.</i>		BOTTINO	03A	PR D67 063519	A. Bottino, N. Fornengo, S. Scopel	
CAO	12A	PL B710 665	J. Cao <i>et al.</i>		CHATTOPADH...	03	PR D68 035005	U. Chattopadhyay, A. Corsetti, P. Nath	
CHATRCHYAN	12	PR D85 012004	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ELLIS	03B	ASP 18 395	J. Ellis, K.A. Olive, Y. Santoso	
CHATRCHYAN	12AE	PRL 109 171803	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ELLIS	03C	NP B552 259	J. Ellis <i>et al.</i>	
CHATRCHYAN	12AI	JHEP 1208 110	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ELLIS	03C	PL B565 176	J. Ellis <i>et al.</i>	
CHATRCHYAN	12AL	JHEP 1206 169	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ELLIS	03D	PL B573 162	J. Ellis <i>et al.</i>	
CHATRCHYAN	12AM	JHEP 1208 026	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ELLIS	03E	PR D67 123502	J. Ellis <i>et al.</i>	
CHATRCHYAN	12AT	JHEP 1210 018	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	HEISTER	03C	EPJ C28 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
CHATRCHYAN	12BJ	JHEP 1211 147	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	HEISTER	03G	EPJ C31 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
CHATRCHYAN	12BK	JHEP 1211 172	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	KLAPDOR-K...	03	ASP 18 525	H.V. Klapdor-Kleingrothaus <i>et al.</i>	
CHATRCHYAN	12BO	JHEP 1212 055	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	LAHANAS	03	PL B568 55	A. Lahanas, D. Nanopoulos	
CHATRCHYAN	12L	PL B713 408	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	TAKEDA	03	PL B572 145	A. Takeda <i>et al.</i>	
DAW	12	ASP 35 397	E. Daw <i>et al.</i>	(DRIFT-III Collab.)	ABRAMS	02	PR D66 122003	D. Abrams <i>et al.</i>	(CDMS Collab.)
DREINER	12A	EPL 99 61001	H.K. Dreiner, M. Kramer, J. Tattersall	(BONN+)	ACHARD	02	PL B524 65	P. Achard <i>et al.</i>	(L3 Collab.)
ELLIS	12B	EPJ C72 2005	J. Ellis, K. Olive		ACOSTA	02H	PRL 89 281801	D. Acosta <i>et al.</i>	(CDF Collab.)
FELIZARDO	12	PRL 108 201302	M. Felizardo <i>et al.</i>	(SIMPLE Collab.)	ANGLOHER	02	ASP 18 43	G. Angloher <i>et al.</i>	(CREST Collab.)
FENG	12B	PR D85 075007	J. Feng, K. Matchev,						

ABREU	01G	PL B503 34	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	01	EPJ C19 397	M. Acciari <i>et al.</i>	(L3 Collab.)
BALTZ	01	PRL 86 5004	E. Baltz, P. Gondolo	
BARATE	01	PL B499 67	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	01B	EPJ C19 415	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARGER	01C	PL B518 117	V. Barger, C. Kao	
BAUDIS	01	PR D63 022001	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
BENOIT	01	PL B513 15	A. Benoit <i>et al.</i>	(EDELWEISS Collab.)
BERNABEI	01	PL B509 197	R. Bernabei <i>et al.</i>	(ALEPH Collab.)
BOTTINO	01	PR D63 125003	A. Bottino <i>et al.</i>	(DAMA Collab.)
CORSETTI	01	PR D64 125010	A. Corsetti, P. Nath	
ELLIS	01B	PL B510 236	J. Ellis <i>et al.</i>	
ELLIS	01C	PR D63 065016	J. Ellis, A. Ferstl, K.A. Olive	
GOMEZ	01	PL B512 252	M.E. Gomez, J.D. Vergados	
LAHANAS	01	PL B518 94	A. Lahanas, D.V. Nanopoulos, V. Spanos	
ABBIENDI	00	EPJ C12 1	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	00G	EPJ C14 51	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	00H	EPJ C14 187	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
Also		EPJ C16 707 (errata.)	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI,G	00D	EPJ C18 253	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	00J	PL B479 129	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00Q	PL B478 65	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00T	PL B485 95	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00U	PL B487 36	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00V	EPJ C16 211	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00W	PL B489 38	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00Z	EPJ C17 53	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABUSADI	00	PRL 84 5699	R. Abusadi <i>et al.</i>	(CDMS Collab.)
ACCIARRI	00C	EPJ C16 1	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	00D	PL B472 420	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCOMANDO	00	NP B585 124	E. Accomando <i>et al.</i>	
BERNABEI	00	PL B480 23	R. Bernabei <i>et al.</i>	(DAMA Collab.)
BERNABEI	00C	EPJ C18 283	R. Bernabei <i>et al.</i>	(DAMA Collab.)
BERNABEI	00D	NJP 2 15	R. Bernabei <i>et al.</i>	(DAMA Collab.)
BOEHM	00B	PR D62 035012	C. Boehm, A. Djouadi, M. Drees	
ELLIS	00	PR D62 075010	J. Ellis <i>et al.</i>	
FENG	00	PL B482 388	J.L. Feng, K.T. Matchev, F. Wilczek	
LEP	00	CERN-EP-2000-016	LEP Collabs.	(ALEPH, DELPHI, L3, OPAL, SLD+)
MORALES	00	PL B489 268	A. Morales <i>et al.</i>	(IGEX Collab.)
PDG	00	EPJ C15 1	D.E. Groom <i>et al.</i>	(PDG Collab.)
SPOONER	00	PL B473 330	N.J.C. Spooner <i>et al.</i>	(UK Dark Matter Col.)
ACCIARRI	99H	PL B456 283	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	99I	PL B459 354	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	99R	PL B470 268	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	99W	PL B471 280	M. Acciari <i>et al.</i>	(L3 Collab.)
AMBROSIO	99	PR D60 082002	M. Ambrosio <i>et al.</i>	(Macro Collab.)
BAUDIS	99	PR D59 022001	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
BELLI	99C	NP B563 97	P. Belli <i>et al.</i>	(DAMA Collab.)
OOTANI	99	PL B461 371	W. Ootani <i>et al.</i>	
ABREU	98P	PL B444 491	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	98F	EPJ C4 207	M. Acciari <i>et al.</i>	(L3 Collab.)
ACKERSTAFF	98P	PL B433 195	K. Ackersstaff <i>et al.</i>	(OPAL Collab.)
BARATE	98K	PL B433 176	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	98S	EPJ C4 433	R. Barate <i>et al.</i>	(ALEPH Collab.)
BERNABEI	98C	PL B436 379	R. Bernabei <i>et al.</i>	(DAMA Collab.)
ELLIS	98	PR D58 095002	J. Ellis <i>et al.</i>	
ELLIS	98B	PL B444 367	J. Ellis, T. Falk, K. Olive	
PDG	98	EPJ C3 1	C. Caso <i>et al.</i>	(PDG Collab.)
BAER	97	PR D57 567	H. Baer, M. Brhlik	
BERNABEI	97	ASP 7 73	R. Bernabei <i>et al.</i>	(DAMA Collab.)
EDSJO	97	PR D56 1879	J. Edsjo, P. Gondolo	
ARNOWITT	96	PR D54 2374	R. Arnowitt, P. Nath	
BAER	96	PR D53 597	H. Baer, M. Brhlik	
BERGSTROM	96	ASP 5 263	L. Bergstrom, P. Gondolo	
LEWIN	96	ASP 6 87	J.D. Lewin, P.F. Smith	
BEREZINSKY	95	ASP 5 1	V. Berezinsky <i>et al.</i>	
FALK	95	PL B354 99	T. Falk, K.A. Olive, M. Srednicki	(MINN, UCSB)
LOSECCO	95	PL B342 392	J.M. LoSecco	(NDAM)
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)
DREES	93	PR D47 376	M. Drees, M.M. Nojiri	(DESY, SLAC)
DREES	93B	PL D48 3483	M. Drees, M.M. Nojiri	
FALK	93	PL B318 354	T. Falk <i>et al.</i>	(UCB, UCSB, MINN)
KELLEY	93	PR D47 2461	S. Kelley <i>et al.</i>	(TAMU, ALAH)
MIZUTA	93	PL B298 120	S. Mizuta, M. Yamaguchi	(TOHO)
MORI	93	PR D48 5505	M. Mori <i>et al.</i>	(KEK, NIIG, TOKY, TOKA+)
BOTTINO	92	MPL A7 733	A. Bottino <i>et al.</i>	(TORI, ZARA)
Also		PL B265 57	A. Bottino <i>et al.</i>	(TORI, INFN)
DECAMP	92	PRPL 216 253	D. Decamp <i>et al.</i>	(ALEPH Collab.)
LOPEZ	92	NP B370 445	J.L. Lopez, D.V. Nanopoulos, K.J. Yuan	(TAMU)
MCDONALD	92	PL B283 80	J.L. McDonald, K.A. Olive, M. Srednicki	(LISB+)
ABREU	91F	NP B367 511	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ALEXANDER	91F	ZPHY C52 175	G. Alexander <i>et al.</i>	(OPAL Collab.)
BOTTINO	91	PL B265 57	A. Bottino <i>et al.</i>	(TORI, INFN)
GELMINI	91	NP B351 623	G.B. Gelmini, P. Gondolo, E. Roulet	(UCLA, TRST)
GRIEST	91	PR D43 3191	K. Griest, D. Seckel	(CHIC, FNAL)
KAMIONKOW	91	PR D44 3021	M. Kamionkowski	(Kamionkowski Collab.)
MORI	91B	PL B270 89	M. Mori <i>et al.</i>	(KEK)
NOJIRI	91	PL B261 76	M.M. Nojiri	(MINN, UCSB)
OLIVE	91	NP B355 208	K.A. Olive, M. Srednicki	(CERN)
ROSZKOWSKI	91	PL B262 59	L. Roszkowski	(UCB+)
GRIEST	90	PR D41 3565	K. Griest, M. Kamionkowski, M.S. Turner	
BARBIERI	89C	NP B313 725	R. Barbieri, M. Frigeni, G. Giudice	
OLIVE	89	PL B230 78	K.A. Olive, M. Srednicki	(MINN, UCSB)
ELLIS	88D	NP B307 883	J. Ellis, R. Flores	
GRIEST	88B	PR D38 2357	K. Griest	
OLIVE	88	PL B205 953	K.A. Olive, M. Srednicki	(MINN, UCSB)
SREDNICKI	88	NP B310 693	M. Srednicki, R. Watkins, K.A. Olive	(MINN, UCSB)
ELLIS	84	NP B238 453	J. Ellis <i>et al.</i>	(CERN)
GOLDBERG	83	PRL 50 1419	H. Goldberg	(NEAS)
KRAUSS	83	NP B227 556	L.M. Krauss	(HARV)
VYSOTSKI	83	SJNP P7 948	M.I. Vysotsky	(ITEP)

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**Technicolor****DYNAMICAL ELECTROWEAK SYMMETRY BREAKING: IMPLICATIONS OF THE  $H^0$** 

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**1. Introduction and Phenomenology**

In theories of dynamical electroweak symmetry breaking, the electroweak interactions are broken to electromagnetism by the vacuum expectation value of a composite operator, typically a fermion bilinear. In these theories, the longitudinal components of the massive weak bosons are identified with composite Nambu-Goldstone bosons arising from dynamical symmetry breaking in a strongly-coupled extension of the standard model. Viable theories of dynamical electroweak symmetry breaking must also explain (or at least accommodate) the presence of an additional composite scalar state to be identified with the  $H^0$  scalar boson [1,2] – a state unlike any other observed to date.

Theories of dynamical electroweak symmetry breaking can be classified by the nature of the composite singlet state to be associated with the  $H^0$ , and the corresponding dimensional scales  $f$ , the analog of the pion decay-constant in QCD, and  $\Lambda$ , the scale of the underlying strong dynamics.<sup>1</sup> Of particular importance is the ratio  $v/f$ , where  $v^2 = 1/(\sqrt{2}G_F) \approx (246 \text{ GeV})^2$ , since this ratio measures the expected size of the deviations of the couplings of a composite Higgs boson from those expected in the standard model. The basic possibilities, and the additional states that they predict, are described below.

**1.1 Technicolor,  $v/f \simeq 1$ ,  $\Lambda \simeq 1 \text{ TeV}$ :**

Technicolor models [8–10] incorporate a new asymptotically free gauge theory (“technicolor”) and additional massless fermions (“technifermions”) transforming under a vectorial representation of the gauge group). The global chiral symmetry of the fermions is spontaneously broken by the formation of a technifermion condensate, just as the approximate chiral symmetry in QCD is broken down to isospin by the formation of a quark condensate. The  $SU(2)_W \times U(1)_Y$  interactions are embedded in the global technifermion chiral symmetries in such a way that the only unbroken gauge symmetry after chiral symmetry breaking is  $U(1)_{em}$ .<sup>2</sup> These theories naturally provide the Nambu-Goldstone bosons “eaten” by the  $W$  and  $Z$  boson. There would also typically be additional heavy states (e.g. vector mesons, analogous to the  $\rho$  and  $\omega$  mesons in QCD) with TeV masses [14,15], and the  $WW$  and  $ZZ$  scattering amplitudes would be expected to be strong at energies of order 1 TeV.

<sup>1</sup> In a strongly interacting theory “Naive Dimensional Analysis” [3,4] implies that, in the absence of fine-tuning,  $\Lambda \simeq g^* f$  where  $g^* \simeq 4\pi$  is the typical size of a strong coupling in the low-energy theory [5,6]. This estimate is modified in the presence of multiple flavors or colors [7].

<sup>2</sup> For a review of technicolor models, see [11–13].

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There are various possibilities for the scalar  $H^0$  in technicolor models, as described below.<sup>3</sup> In all of these cases, however, to the extent that the  $H^0$  has couplings consistent with those of the standard model, these theories are very highly constrained.

- a)  **$H^0$  as a singlet scalar resonance:** The strongly-interacting fermions which make up the Nambu-Goldstone bosons eaten by the weak bosons would naturally be expected to also form an isoscalar neutral bound state, analogous to the  $\sigma$  particle expected in pion-scattering in QCD [16]. However, in this case, there is no symmetry protecting the mass of such a particle – which would therefore generically be of order the energy scale of the underlying strong dynamics  $\Lambda$ . In the simplest theories of this kind – those with a global  $SU(2)_L \times SU(2)_R$  chiral symmetry which is spontaneously broken to  $SU(2)_V$  – the natural dynamical scale  $\Lambda$  would be of order a TeV, resulting in a particle too heavy and broad to be identified with the  $H^0$ . The scale of the underlying interactions could naturally be smaller than 1 TeV if the global symmetries of the theory are larger than  $SU(2)_L \times SU(2)_R$ , but in this case there would be additional (pseudo-)Nambu-Goldstone bosons (more on this below). A theory of this kind would only be viable, therefore, if some choice of the parameters of the high energy theory could give rise to sufficiently light state without the appearance of additional particles that should have already been observed. Furthermore, while a particle with these quantum numbers could have Higgs-like couplings to any electrically neutral spin-zero state made of quarks, leptons, or gauge-bosons, there is no symmetry insuring that the coupling strengths of such a composite singlet scalar state would be precisely the same as those of the standard model Higgs [17].
- b)  **$H^0$  as a dilaton:** It is possible that the underlying strong dynamics is approximately scale-invariant, as inspired by theories of “walking technicolor” [18–22], and that both the scale and electroweak symmetries are spontaneously broken at the TeV energy scale [23]. In this case, due to the spontaneous breaking of approximate scale invariance, one might expect a corresponding (pseudo-) Nambu-Goldstone boson [19] with a mass less than a TeV, the dilaton.<sup>4</sup> A dilaton couples to the trace of the energy momentum tensor, which leads to a similar pattern of two-body couplings as the couplings of the standard model Higgs boson [28–30]. Scale-invariance is a space-time symmetry, however, and is unrelated to the global symmetries that we can identify with the electroweak group. Therefore the

<sup>3</sup> In these models, the self-coupling of the  $H^0$  scalar is not related to its mass, as it is in the SM – though there are currently no experimental constraints on this coupling.

<sup>4</sup> Even in this case, however, a dilaton associated with electroweak symmetry breaking will likely not *generically* be as light as the  $H^0$  [24–27].

decay-constants associated with the breaking of the scale and electroweak symmetries will not, in general, be the same.<sup>5</sup> In other words, if there are no large anomalous dimensions associated with the  $W$ - and  $Z$ -bosons or the top- or bottom-quarks, the ratios of the couplings of the dilaton to these particles would be the same as the ratios of the same couplings for the standard model Higgs boson, but the overall strength of the dilaton couplings would be expected to be different [31,32]. Furthermore, the couplings of the dilaton to gluon- and photon-pairs can be related to the beta functions of the corresponding gauge interactions in the underlying high-energy theory, and will not in general yield couplings with the exactly the same strengths as the standard model [33,34].

- c)  **$H^0$  as a singlet Pseudo-Nambu-Goldstone Boson:** If the global symmetries of the technicolor theory are larger than  $SU(2)_L \times SU(2)_R$ , there can be extra singlet (pseudo-) Nambu-Goldstone bosons which could be identified with the  $H^0$ . In this case, however, the coupling strength of the singlet state to  $WW$  and  $ZZ$  pairs would be comparable to the couplings to gluon and photon pairs, and these would all arise from loop-level couplings in the underlying technicolor theory [35]. This pattern of couplings is not supported by the data.

### 1.2 The Higgs doublet as a pseudo-Nambu-Goldstone Boson, $v/f < 1$ , $\Lambda > 1$ TeV:

In technicolor models, the symmetry-breaking properties of the underlying strong dynamics necessarily breaks the electroweak gauge symmetries. An alternative possibility is that the underlying strong dynamics itself does not break the electroweak interactions, and that the entire quartet of bosons in the Higgs doublet (including the state associated with the  $H^0$ ) are composite (pseudo-) Nambu-Goldstone particles [36,37]. In this case, the underlying dynamics can occur at energies larger than 1 TeV and additional interactions with the top-quark mass generating sector (and possibly with additional weakly-coupled gauge bosons) cause the vacuum energy to be minimized when the composite Higgs doublet gains a vacuum expectation value [38,39]. In these theories, the couplings of the remaining singlet scalar state would naturally be equal to that of the standard model Higgs boson up to corrections of order  $(v/f)^2$  and, therefore, constraints on the size of deviations of the  $H^0$  couplings from that of the standard model Higgs give rise to lower bounds on the scales  $f$  and  $\Lambda$ .<sup>6</sup>

<sup>5</sup> If both the electroweak symmetry and the approximate scale symmetry are broken only by electroweak doublet condensate(s), then the decay-constants for scale and electroweak symmetry breaking may be approximately equal – differing only by terms formally proportional to the amount of explicit scale-symmetry breaking.

<sup>6</sup> In these models  $v/f$  is an adjustable parameter, and in the limit  $v/f \rightarrow 1$  they reduce, essentially, to the technicolor models discussed in the previous subsection. Our discussion here is



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The electroweak gauge interactions, as well as the interactions responsible for the top-quark mass, explicitly break the chiral symmetries of the composite Higgs model, and lead generically to sizable corrections to the mass-squared of the Higgs-doublet – the so-called “Little Hierarchy Problem” [40]. “Little Higgs” theories [41–44] are examples of composite Higgs models in which the (collective) symmetry-breaking structure is selected so as to suppress these contributions to the Higgs mass-squared.

Composite Higgs models typically require a larger global symmetry of the underlying theory, and hence additional relatively light (compared to  $\Lambda$ ) scalar particles, extra electroweak vector bosons (e.g. an additional  $SU(2) \times U(1)$  gauge group), and vector-like partners of the top-quark of charge  $+2/3$  and possibly also  $+5/3$  [45]. Finally, in addition to these states, one would expect the underlying dynamics to yield additional scalar and vector resonances with masses of order  $\Lambda$ . If the theory respects a custodial symmetry [46], the couplings of these additional states to the electroweak and Higgs boson will be related – and, for example, one might expect a charged vector resonance to have similar branching ratios to  $WZ$  and  $WH$ . Different composite Higgs models utilize different mechanisms for arranging for the hierarchy of scales  $v < f$  and arranging for a scalar Higgs self-coupling small enough to produce an  $H^0$  of mass of order 125 GeV, for a review see [48].

### 1.3 Top-Condensate, Top-Color, Top-Seesaw and related theories, $v/f < 1$ , $\Lambda > 1$ TeV:

A final alternative is to consider a strongly interacting theory with a high (compared to a TeV) underlying dynamical scale that *would* naturally break the electroweak interactions, but whose strength is adjusted (“fine-tuned”) to produce electroweak symmetry breaking at 1 TeV. This alternative is possible if the electroweak (quantum) phase transition is continuous (second order) in the strength of the strong dynamics [47]. If the fine tuning can be achieved, the underlying strong interactions will produce a light composite Higgs bound state with couplings equal to that of the standard model Higgs boson up to corrections of order  $(1 \text{ TeV}/\Lambda)^2$ . As in theories in which electroweak symmetry breaking occurs through vacuum alignment, therefore, constraints on the size of deviations of the  $H^0$  couplings from that of the standard model Higgs give rise to lower bounds on the scale  $\Lambda$ . Formally, in the limit  $\Lambda \rightarrow \infty$  (a limit which requires arbitrarily fine adjustment of the strength of the high-energy interactions), these theories are equivalent to a theory with a fundamental Higgs boson – and the fine adjustment of the coupling strength is a manifestation of the hierarchy problem of theories with a fundamental scalar particle.

In many of these theories the top-quark itself interacts strongly (at high energies), potentially through an extended consistent with that given there, since we expect corrections to the SM Higgs couplings to be large for  $v/f \simeq 1$ .

color gauge sector [49–53]. In these theories, top-quark condensation (or the condensation of an admixture of the top with additional vector-like quarks) is responsible for electroweak symmetry breaking, and the  $H^0$  is identified with a bound state involving the third generation of quarks. These theories typically include an extra set of massive color-octet vector bosons (top-gluons), and an extra  $U(1)$  interaction (giving rise to a top-color  $Z'$ ) which couple preferentially to the third generation and whose masses define the scale  $\Lambda$  of the underlying physics.

### 1.4 Flavor

In addition to the electroweak symmetry breaking dynamics described above, which gives rise to the masses of the  $W$  and  $Z$  particles, additional interactions must be introduced to produce the masses of the standard model fermions. Two general avenues have been suggested for these new interactions. In one case, e.g. “extended technicolor” (ETC) theories [54,55], the gauge interactions in the underlying strongly interacting theory are extended to incorporate flavor. This extended gauge symmetry is broken down (possibly sequentially, at several different mass scales) to the residual strongly-interacting interaction responsible for electroweak symmetry breaking. The massive gauge-bosons corresponding to the broken symmetries then mediate interactions between mass operators for the quarks/leptons and the corresponding bilinears of the strongly-interacting fermions, giving rise to the masses of the ordinary fermions after electroweak symmetry breaking. An alternative proposal, “partial compositeness” [56], the additional interactions giving rise to mixing between the ordinary quarks and leptons and massive composite fermions in the strongly-interacting underlying theory. Theories incorporating partial compositeness include additional vector-like partners of the ordinary quarks and leptons, typically with masses of order a TeV or less.

In both cases, the effects of these flavor interactions on the electroweak properties of the ordinary quarks and leptons are likely to be most pronounced in the third generation of fermions.<sup>7</sup> The additional particles present, especially the additional scalars, often couple more strongly to heavier fermions.

Moreover, since the flavor interactions must give rise to quark mixing, we expect that a generic theory of this kind could give rise to large flavor-changing neutral-currents [55]. In ETC theories, these constraints are typically somewhat relaxed if the theory incorporates approximate generational flavor symmetries [57], the theory “walks” [18–22], or if  $\Lambda > 1$  TeV [58]. In theories of partial compositeness, the masses of the ordinary fermions depend on the scaling-dimension of the operators corresponding to the composite fermions with which they mix. This leads to a new mechanism for generating the

<sup>7</sup> Indeed, from this point of view, the vector-like partners of the top-quark in top-seesaw and little Higgs models can be viewed as incorporating partial compositeness to explain the origin of the top quark’s large mass.

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mass-hierarchy of the observed quarks and leptons that, potentially, ameliorates flavor-changing neutral current problems and can provide new contributions to the composite Higgs potential which allows for  $v/f < 1$  [59–63].

Alternatively, one can assume that the underlying flavor dynamics respects flavor symmetries (“minimal” [64,65] or “next-to-minimal” [66] flavor violation) which suppress flavor-changing neutral currents in the two light generations. Additional considerations apply when extending these considerations to potential explanation of neutrino masses (see, for example, [67,68]).

Since the underlying high-energy dynamics in these theories are strongly coupled, there are no reliable calculation techniques that can be applied to analyze their properties. Instead, most phenomenological studies depend on the construction of a “low-energy” effective theory describing additional scalar, fermion, or vector boson degrees of freedom, which incorporates the relevant symmetries and, when available, dynamical principles. In some cases, motivated by the AdS/CFT correspondence [69], the strongly-interacting theories described above have been investigated by analyzing a dual compactified five-dimensional gauge theory. In these cases, the AdS/CFT “dictionary” is used to map the features of the underlying strongly coupled high-energy dynamics onto the low-energy weakly coupled dual theory [70].

More recently, progress has been made in investigating strongly-coupled models using lattice gauge theory [71–73]. These calculations offer the prospect of establishing which strongly coupled theories of electroweak symmetry breaking have a particle with properties consistent with those observed for the  $H^0$  – and for establishing concrete predictions for these theories at the LHC [74].

## 2. Experimental Searches

As discussed above, the extent to which the couplings of the  $H^0$  conform to the expectations for a standard model Higgs boson constrains the viability of each of these models. Measurements of the  $H^0$  couplings, and their interpretation in terms of effective field theory, are summarized in the  $H^0$  review in this volume. In what follows, we will focus on searches for the additional particles that might be expected to accompany the singlet scalar: extra scalars, fermions, and vector bosons. In some cases, detailed model-specific searches have been made for the particles described above (though generally not yet taking account of the demonstrated existence of the  $H^0$  boson).

In most cases, however, generic searches (e.g. for extra  $W'$  or  $Z'$  particles, extra scalars in the context of multi-Higgs models, or for fourth-generation quarks) are quoted that can be used – when appropriately translated – to derive bounds on a specific model of interest.

The mass scale of the new particles implied by the interpretations of the low mass of  $H^0$  discussed above, and existing studies from the Tevatron and lower-energy colliders, suggests that only the Large Hadron Collider has any real sensitivity.

A number of analyses already carried out by ATLAS and CMS use relevant final states and might have been expected to observe a deviation from standard model expectations – in no case so far has any such deviation been reported. The detailed implications of these searches in various model frameworks are described below.

Except where otherwise noted, all limits in this section are quoted at a confidence level of 95%. The ATLAS searches have analyzed  $20.3 \text{ fb}^{-1}$  of data recorded at the LHC with  $\sqrt{s}=8 \text{ TeV}$ , and the CMS analyses are based on the data collected at  $\sqrt{s} = 8 \text{ TeV}$  in 2012 with an integrated luminosity of  $19.7 \text{ fb}^{-1}$ .

### 2.1 Searches for $Z'$ or $W'$ Bosons

Massive vector bosons or particles with similar decay channels would be expected to arise in Little Higgs theories, in theories of Technicolor, or models involving a dilaton, adjusted to produce a light Higgs boson, consistent with the observed  $H^0$ . These particles would be expected to decay to pairs of vector bosons, to third generation quarks, or to leptons. The generic searches for  $W'$  and  $Z'$  vector bosons listed below can, therefore, be used to constrain models incorporating a composite Higgs-like boson.

#### $Z' \rightarrow \ell\ell$ :

ATLAS [76] and CMS [77] have both searched for  $Z'$  production with  $Z' \rightarrow ee$  or  $\mu\mu$ . The main backgrounds to these analyses arise from Drell-Yan,  $t\bar{t}$ , and diboson production and are estimated using Monte Carlo simulation, with the cross sections scaled by next-to-next-to-leading-order  $k$ -factors. Instrumental backgrounds from QCD multijet and  $W$ +jet events are estimated using control data samples. No deviation from the standard model prediction is seen in the dielectron and dimuon invariant mass spectra, by either the ATLAS or the CMS analysis, and lower limits on possible  $Z'$  boson masses are set. The dielectron channel has higher sensitivity due to the superior mass resolution compared to the dimuon channel. A  $Z'_{\text{SSM}}$  with couplings equal to the standard model  $Z$  (a “sequential standard model”  $Z'$ ) and a mass below 2.79 TeV is excluded by ATLAS, while CMS sets a lower mass limit of 2.90 TeV. The ATLAS analysis rules out various  $E6$ -motivated bosons ( $Z'_\psi, Z'_\chi$ ) and  $Z^*$  with masses lower than 2.51, 2.62 and 2.85 TeV, while a  $Z'_\psi$  with a mass below 2.57 TeV is excluded by CMS. The experiments also place limits on the parameters of extra dimension models and in the case of ATLAS on the parameters of a minimal walking technicolor model [18–22], consistent with a 125 GeV Higgs boson.

In addition, both experiments have also searched for  $Z'$  decaying to a ditau final state [78,79]. While less sensitive than dielectron or dimuon final states, an excess in  $\tau^+\tau^-$  could have interesting implications for models in which lepton universality is not a necessary requirement and enhanced couplings to the third generation are allowed. This analysis leads to lower limits on the mass of a  $Z'_{\text{SSM}}$  of 2.0 and 1.3 TeV from ATLAS and CMS respectively.

$Z' \rightarrow q\bar{q}$ :

The ability to relatively cleanly select  $t\bar{t}$  pairs at the LHC together with the existence of enhanced couplings to the third generation in many models makes it worthwhile to search for new particles decaying in this channel. Both ATLAS [80] and CMS [81] have carried out searches for new particles decaying into  $t\bar{t}$ . ATLAS focuses on the lepton plus jets final state, where the top quark pair decays as  $t\bar{t} \rightarrow WbWb$  with one  $W$  boson decaying leptonically and the other hadronically; CMS uses final states where both, one or neither  $W$  decays leptonically and then combines the results. The  $t\bar{t}$  invariant mass spectrum is analyzed for any excess, and no evidence for any resonance is seen. ATLAS excludes a narrow ( $\Gamma/m = 1.2\%$ ) leptophobic top-color  $Z'$  boson with a mass below 1.8 TeV; upper limits are set on the cross section times branching ratio for a broad color octet resonance with  $\Gamma/m = 15\%$  decaying to  $t\bar{t}$  which range from 4.8 pb for  $m = 0.4$  TeV to 0.09 pb for  $m = 3.0$  TeV. CMS sets limits on a narrow ( $\Gamma/m = 1.2\%$ )  $Z'$  boson decaying to  $t\bar{t}$  of 2.4 TeV and on a wide resonance (10% width) of 2.8 TeV. In the Randall-Sundrum model, KK gravitons ( $g_{KK}$ ) with masses below 2.2 TeV are excluded by ATLAS and (for a different set of model parameters) below 2.7 TeV by CMS.

Both ATLAS [82] and CMS [83] have also searched for resonances decaying into  $q\bar{q}$ ,  $qg$  or  $gg$  using the dijet invariant mass spectrum. Model-independent upper limits on cross sections are set; ATLAS excludes color-octet scalars below 2.72 TeV,  $W'$  bosons below 2.45 TeV and chiral  $W^*$  bosons below 1.75 TeV. CMS is able to exclude  $W'$  bosons below 1.9 TeV or between 2.0 and 2.2 TeV;  $Z'$  bosons below 1.7 TeV; and  $g_{KK}$  gravitons below 1.6 TeV. Searches are also carried out for wide resonances, assuming  $\Gamma/m$  up to 30%, and exclude axigluons and colorons with mass below 3.6 TeV, and color-octet scalars with mass below 2.5 TeV.

 $W' \rightarrow \ell\nu$ :

Both LHC experiments have also searched for massive charged vector bosons. ATLAS [85] searched for a heavy  $W'$  decaying to  $e\nu$  or  $\mu\nu$  and find no excess over the standard model expectation. A sequential standard model  $W'$  (assuming zero branching ratio to  $WZ$ ) with mass less than 3.24 TeV is excluded, and excited chiral bosons  $W^*$  excluded up to 3.21 TeV.

CMS [86] has carried out a complementary search in the  $\tau\nu$  final state. As noted above, such searches place interesting limits on models with enhanced couplings to the third generation. No excess is observed and limits between 2.0 and 2.7 TeV are set on the mass of a  $W'$  decaying preferentially to the third generation; a  $W'$  with universal fermion couplings is also excluded for masses less than 2.7 TeV.

 $W' \rightarrow t\bar{b}$ :

Heavy new gauge bosons can couple to left-handed fermions like the  $W$  boson or to right-handed fermions.  $W'$  bosons that couple only to right-handed fermions may not have leptonic decay modes, depending on the mass of the right-handed

neutrino. For these  $W'$  bosons, the  $t\bar{b}$  decay mode is especially important because it is the hadronic decay mode with the best signal-to-background.

ATLAS has searched for  $W'$  bosons in the  $t\bar{b}$  final state both for leptonic [87] and hadronic [88] decays of the top. No significant deviations from the standard model are seen in either analysis and limits are set on the  $W' \rightarrow t\bar{b}$  cross section times branching ratio and on the  $W'$  effective couplings.  $W'$  bosons with purely left-handed (right-handed) couplings to fermions are excluded for masses below 1.70 (1.92) TeV.

**2.2 Searches for Resonances decaying to Vector Bosons and/or Higgs Bosons** $X \rightarrow WW, WZ, ZZ$ :

Both experiments have used the data collected at  $\sqrt{s} = 8$  TeV to search for resonances decaying to pairs of bosons.

ATLAS [89] and CMS [90] have both looked for a resonant state (such as a  $W'$ ) decaying to  $WZ$  in the fully-leptonic channel,  $\ell\nu\ell'\ell'$  (where  $\ell, \ell' = e, \mu$ ). The  $WZ$  invariant mass distribution reconstructed from the observed lepton momenta missing transverse energy. The backgrounds arise mainly from standard model  $WZ$ ,  $ZZ$  and  $t\bar{t} + W/Z$  production. No significant deviation from the standard model prediction is observed by either experiment. A  $W'$  with mass less than 1.55 (1.52) TeV is excluded by CMS (ATLAS); ATLAS also sets limits on the production cross section for heavy vector triplet particles, and CMS sets limits on the production of low-scale technimesons  $\rho_{TC}$  from the reconstructed  $WZ$  mass spectrum and cross section.

ATLAS [91,92] and CMS [93] have also searched for narrow resonances decaying to  $WW$ ,  $WZ$  or  $ZZ$  in  $\ell\nu jj$  and  $\ell\ell jj$  final states (where one boson decays leptonically and the other to jets). No deviation from the standard model is seen by either experiment; resonance masses below 1.59 TeV for an extended gauge model  $W'$  are excluded by ATLAS. CMS interprets their results in terms of Randall-Sundrum  $g_{KK}$  production but also presents model-independent cross section limits that can be used to constrain other models.

Searches have also been conducted in fully hadronic final states. ATLAS [94] and CMS [95] have searched for massive resonance in dijet systems with one or both jets identified as a  $W$  or a  $Z$  boson using jet-substructure techniques. ATLAS observes a small excess (less than three standard deviations) around 2 TeV in the  $WZ$  channel but otherwise no deviations from the standard model are seen. Limits are set by both experiments on the production cross section times branching ratio for new heavy  $W'$  decaying to  $WZ$  and for  $g_{KK}$  gravitons decaying to  $WW$  or  $ZZ$ . CMS also sets limits on the production of particles decaying to  $qW$  and  $qZ$ .

 $X \rightarrow W/Z + H^0$  and  $X \rightarrow H^0 H^0$ :

With the existence and decay properties of the Higgs boson established, and the significant datasets now available, it is

## Searches Particle Listings

### Technicolor

possible to use searches for anomalous production of the Higgs as a potential signature for new physics. ATLAS [96] and CMS [97,98] have both searched in the data collected at  $\sqrt{s} = 8$  TeV for new particles decaying to a vector boson plus a Higgs boson, where the vector boson decays leptonically (ATLAS) or hadronically (CMS) and the Higgs boson to  $b\bar{b}$  (both experiments),  $WW$  or  $\tau^+\tau^-$  (CMS). No deviation from the standard model is seen in any of these final states and limits can be placed on the allowed production cross section times branching ratio for resonances between 0.8 and 2.5 TeV, on the parameters of a Minimal Walking Technicolor Model and on a heavy vector triplet model.

Both experiments [99,100] have also searched for resonant production of Higgs boson pairs  $X \rightarrow H^0 H^0$  with  $H^0 \rightarrow b\bar{b}$ . No signal is observed and limits are placed on the possible production cross section for any new resonance.

#### $Y \rightarrow W/Z + X$ with $X \rightarrow jj$ :

ATLAS has searched for a dijet resonance [101] with an invariant mass in the range 130 – 300 GeV, produced in association with a  $W$  or a  $Z$  boson. The analysis used  $20.3 \text{ fb}^{-1}$  of data recorded at  $\sqrt{s} = 8$  TeV. The  $W$  or  $Z$  boson is required to decay leptonically ( $\ell = e, \mu$ ). No significant deviation from the standard model prediction is observed and limits are set on the production cross section times branching ratio for a hypothetical technipion produced in association with a  $W$  or  $Z$  boson from the decay of a technirho particle in the context of Low Scale Technicolor models.

### 2.3 Vector-like third generation quarks

Vector-like quarks (VLQ) have non-chiral couplings to  $W$  bosons, i.e. their left- and right-handed components couple in the same way. They therefore have vectorial couplings to  $W$  bosons. Vector-like quarks arise in Little Higgs theories, top-coloron-models, and theories of a composite Higgs boson with partial compositeness. At the LHC, VLQs can be pair produced, via the dominant gluon-gluon fusion. VLQs can also be produced singly by their electroweak effective couplings to a weak boson and a standard model quark. In the following the notation  $T$  quark refers to a vector-like quark with charge  $2/3$  and the notation  $B$  quark refers to a vector-like quark with charge  $-1/3$ .  $T$  quarks can decay to  $bW$ ,  $tZ$ , or  $tH^0$ . Weak isospin singlets are expected to decay to all three final states with (asymptotic) branching fractions of 50%, 25%, 25%, respectively. Weak isospin doublets are expected to decay exclusively to  $tZ$  and to  $tH^0$  [102]. Analogously,  $B$  quarks can decay to  $tW$ ,  $bZ$ , or  $bH^0$ .

### Searches for $T$ quarks that decay to $W$ , $Z$ and $H^0$ bosons

#### $T \rightarrow bW$ :

CMS has searched for pair production of heavy  $T$  quarks that decay exclusively to  $bW$  [103]. The analysis selects events with exactly one charged lepton, assuming that the  $W$  boson from the second  $T$  quark decays hadronically. Under

this hypothesis, a 2-constraint kinematic fit can be performed to reconstruct the mass of the  $T$  quark. The two-dimensional distribution of reconstructed mass vs  $S_T$  is used to test for the signal.  $S_T$  is the scalar sum of the missing  $p_T$  and the transverse momenta of the lepton and the leading four jets. At times the hadronically-decaying  $W$  boson is produced with a large Lorentz boost, leading to the  $W$  decay products merged into a wide single jet also known as a fat jet. Algorithms such as jet pruning [104] are used to resolve the substructure of the fat jets from the decays of the heavy particles. If the mass of the boosted jet is compatible with the  $W$ -boson mass, then this  $W$  boson candidate jet used in the kinematic reconstruction of the  $T$  quark. No excess over standard model backgrounds is observed. This analysis, when combined with the search in the fully hadronic final state [105] excludes new quarks that decay 100% to  $bW$  for masses below 890 GeV [106].

An analogous search has been carried out by ATLAS [109]. It uses the lepton+jets final state with an isolated electron or muon and at least four jets, at least one of which must be tagged as a b-jet. The selection is optimized for  $T$  quark masses above about 400 GeV and requires reconstruction of hadronically decaying  $W$  boson, including those with a high boost leading to merged decay products, and large angular separation between the  $W$  bosons and the b-jets originating from the decay of the heavy  $T$  quark. The analysis focuses on the reconstructed heavy  $T$  quark mass from the hadronic  $W$  candidate and a b-jet. No significant excess of events above standard model expectation is observed. For  $BR(T \rightarrow bW) = 1$ ,  $T$  quark masses lower than 765 GeV are excluded.

#### $T \rightarrow tH^0$ :

ATLAS has performed a search for  $T\bar{T}$  production with  $T \rightarrow tH^0$  [109]. Given the dominant decay mode  $H^0 \rightarrow b\bar{b}$ , these events are characterized by a large number of jets, many of which are b-jets. Thus the event selection requires one isolated electron or muon and at least five jets, two of which must be tagged as b-jets. The data are classified according to their jet-multiplicity (five and six-or-more), b-jet multiplicity (2, 3, and  $\geq 4$ ) and the invariant mass of the two b-tagged jets with lowest  $\Delta R$  between the two b-tagged jets (for  $\geq$  six jet events). The distribution of  $H_T$ , the scalar sum of the lepton and jet  $p_T$ s and the missing  $E_T$ , for each category is used as the discriminant for the final signal and background separation. No excess of events is found. Weak isospin doublet  $T$  quarks are excluded below 855 GeV for  $BR(T \rightarrow tH^0) = 1$ . The CMS search for  $T\bar{T}$  production, with  $T \rightarrow tH^0$  decays have been performed in both lepton+jets, multilepton and all hadronic final states. The lepton+jets analysis [110] emphasizes the presence of large number of b-tagged jets, and combines with other kinematic variables in a Boosted Decision Tree (BDT) for enhancing signal to background discrimination. The multilepton analysis [110] optimized for the presence of b-jets and the large hadronic activity. For  $BR(T \rightarrow Wb) = 1$ , the combined lepton+jets and multilepton analyses lead to a lower limit on  $T$  quark

masses of 706 GeV. A search for  $T \rightarrow tH^0$  in all hadronic decays [111], optimized for a high mass  $T$  quark, and based on identifying boosted top quark jets has been carried out by CMS. This search aims to resolve sub-jets within the fat jet arising from boosted top quark decays, including b-tagging of the sub-jets. A likelihood discriminator is defined based on the distributions of  $H_T$ , and the invariant mass of the two b-jets in the events for signal and background. No excess above background expectations is observed. Assuming 100% BR for  $T \rightarrow tH^0$ , this analysis leads to a lower limit of 745 GeV on the mass of the  $T$  quark.

A CMS search for  $T \rightarrow tH^0$  with  $H^0 \rightarrow \gamma\gamma$  decays has been performed [112]. To identify the Higgs boson produced in the decay of the heavy  $T$  quark, and the subsequent  $H^0 \rightarrow \gamma\gamma$  decay, the analysis focuses on identification of two photons in events with one or more high  $p_T$  lepton+jets or events with no leptons and large hadronic activity. A search for a resonance in the invariant mass distribution of the two photons in events with large hadronic activity defined by the  $H_T$  variable shows no excess above the prediction from standard model processes. The analysis results in exclusion of  $T$  quark masses below 540 GeV.

#### $T \rightarrow tZ$ :

A targeted search by CMS for  $T$  quarks that decay exclusively to  $tZ$  based on an integrated luminosity of  $1.1 \text{ fb}^{-1}$  from pp collisions at  $\sqrt{s} = 7 \text{ TeV}$  [107]. Selected events must have three isolated charged leptons, two of which must be consistent with a leptonic  $Z$ -boson decay. No significant excess was observed.  $T$  quark masses below 485 GeV are excluded. The CMS analysis [110] with combined searches in lepton+jets, dilepton and multilepton final states yields a lower limit on the mass of the  $T$  of 782 GeV. A complementary search has been carried out by ATLAS for new heavy quarks decaying into a  $Z$  boson and a third generation quark [113], with  $T \rightarrow tZ$ . Selected events contain a high transverse momentum  $Z$  boson that decays leptonically, together with two b-jets, which is modified to require at least one b-tagged jet, for events with additional leptons. No significant excess of events above the standard model expectation is observed. For the weak-isospin singlet scenario, a  $T$  quark with mass lower than 655 GeV is excluded, while for a particular weak-isospin doublet scenario, a  $T$  quark with mass lower than 735 GeV is excluded.

The ATLAS experiment has studied the electroweak production of single  $T$  quarks, which is accompanied by a b-jet and a light jet [113]. The initial event selection for this search is very similar to that of  $T\bar{T}$  production with  $T \rightarrow tZ$  decays, and for both the dilepton and trilepton signatures, it requires the presence of an additional energetic jet in the forward region ( $2.5 < |\eta| < 4.5$ ), a characteristic signature of single heavy quark production. An upper limit of 190 fb is obtained for the process  $\sigma(pp \rightarrow T\bar{t}q) \times B(T \rightarrow tZ)$  with a heavy  $T$  quark mass at 700 GeV. For a specific composite Higgs model [114], the  $WTb$  vertex is parameterized by  $\lambda_T$ , which is associated with

the Yukawa coupling in the composite sector and the degree of compositeness of the quarks in the 3<sup>rd</sup> generation. With the current dataset unfortunately the search is not sensitive to  $\lambda_T < 1.5$  nor to any values of  $V_{Tb} < 1$ .

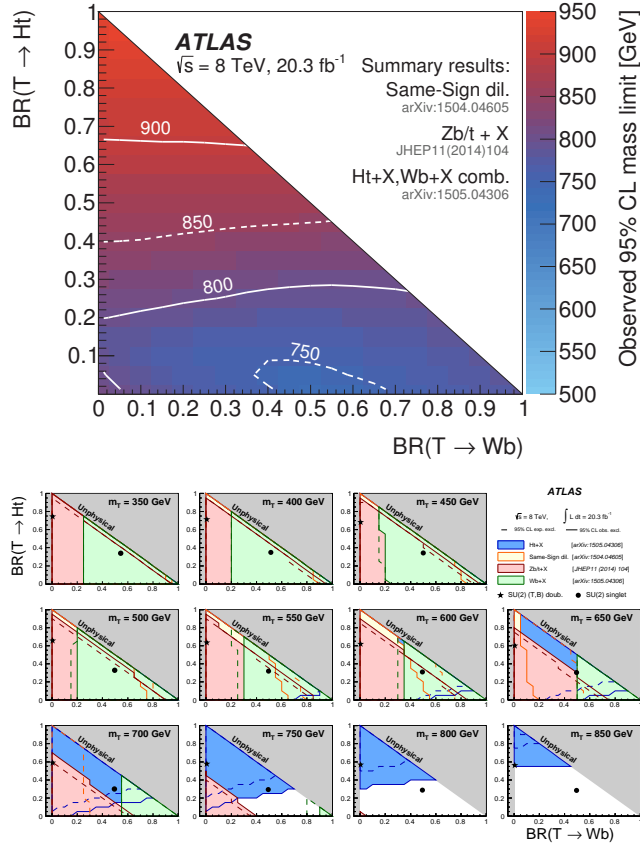
#### Same-Sign dilepton analyses:

Pair-production of  $T$  or  $B$  quarks with their antiparticles can result in events with like-sign leptons, for example if the decay  $T \rightarrow tH \rightarrow bWW^+W^-$  is present, followed by leptonic decays of two same-sign  $W$  bosons. ATLAS and CMS have searched for this final state. The ATLAS search [121] requires two leptons with the same electric charge, at least two jets of which at least one must be tagged as a b-jet, and missing  $p_T$ . ATLAS quotes exclusions of some possible branching fraction combinations depending on the mass of the new quarks.  $T$  quarks that are electroweak singlets are excluded below 590 GeV (assuming branching fractions to the  $W$ ,  $Z$ , and  $H^0$  decay modes arising from a singlet model). For the same-sign lepton signature, the sensitivity is largest for  $T$  quarks that decay exclusively to  $tH^0$ .

#### Combination $T \rightarrow bW/tZ/tH^0$ :

The limits set by ATLAS searches in lepton+jets, dileptons with same-sign charge, and final states with  $Z$  boson have been combined and the results obtained for various combinations of branching fractions for  $T$  quark decays to  $bW$ ,  $tH^0$  and  $tZ$  are shown in Fig. 1. The combined analysis excludes  $T$  quarks that exclusively decay to  $bW/tH^0$  with masses below 765/950 GeV [109], and sets lower  $T$  quarks mass limits that range from 715 to 950 GeV for all possible values of the branching fractions to the three decay modes.

An inclusive search by CMS targeted at heavy  $T$  quarks decaying to any combination of  $bW$ ,  $tZ$ , or  $tH^0$  is described in Ref. [110]. Selected events have at least one isolated charged lepton. Events are categorized according to number and flavour of the leptons, the number of jets, and the presence of hadronic vector boson and top quark decays that are merged into a single jet. The use of jet substructure to identify hadronic decays significantly increases the acceptance for high  $T$  quark masses. No excess above standard model backgrounds is observed. Limits on the pair production cross section of the new quarks are set, combining all event categories, for all combinations of branching fractions into the three final states. For  $T$  quarks that exclusively decay to  $bW/tZ/tH^0$ , masses below 700/782/706 GeV are excluded. Electroweak singlet vector-like  $T$  quarks which decay 50% to  $bW$ , 25% to  $tZ$ , and 25% to  $tH^0$  are excluded for masses below 696 GeV (Fig. 2 top panel). This analysis was the first from CMS to obtain limits on the mass of the  $T$  quark for all possible values of the branching fractions into the three different final states  $bW$ ,  $tZ$  and  $tH$  [110]. A combination [106] of the leptonic inclusive analysis with the targeted  $T \rightarrow tH^0$  decays to all-hadronic final state, and  $T \rightarrow Wb$  decays with all-hadronic and single-lepton final states with emphasis on  $bW$  mass reconstruction, leads to a combined exclusion of  $T$  quarks between 790 and 890 GeV and is shown



**Figure 1:** Observed limits on the mass of the  $T$  quark in the plane of  $BR(T \rightarrow tH^0)$  versus  $BR(T \rightarrow bW)$  from all ATLAS searches for  $TT$  production [109]. Top panel: summary of the most restrictive observed limit on the mass. Contour lines are provided to guide the eye. Bottom panel: Exclusion limits are drawn sequentially for each of the analyses and overlaid (rather than combined). The circle and star symbols denote the default branching ratios for the weak-isospin singlet and doublet cases.

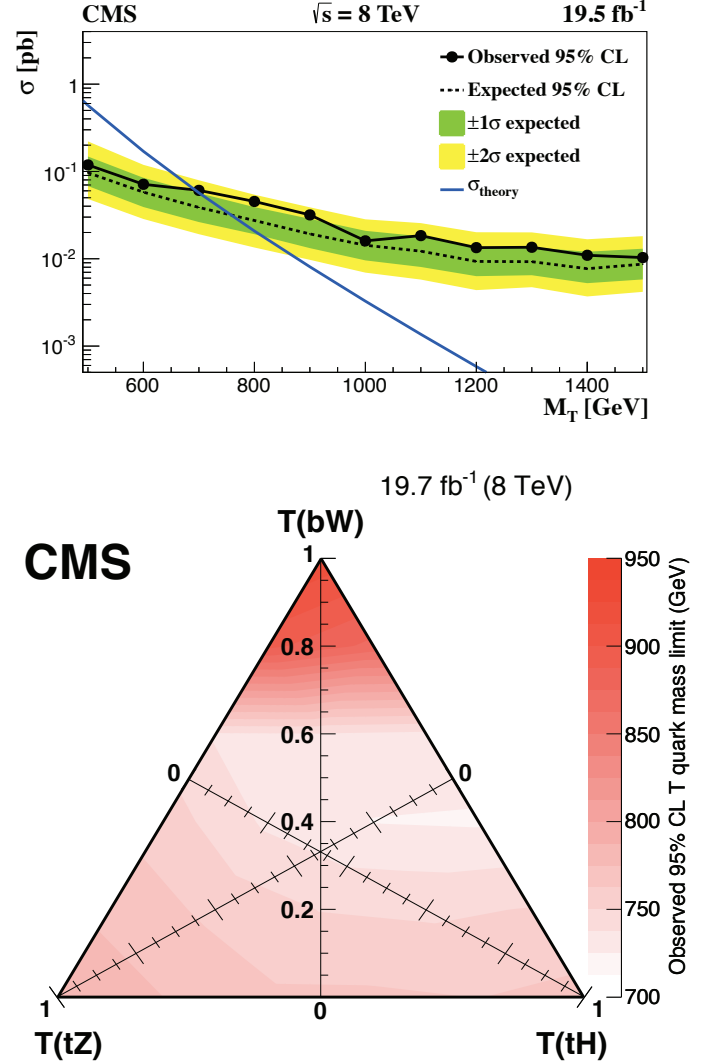
in Fig. 2 (bottom panel). From the combination analyses, any  $T$  quark that exclusively decays to  $bW/tZ/tH$  is required to have masses above 890/830/840 GeV [106].

#### Searches $B$ quarks that decay to $W$ , $Z$ and $H^0$ bosons

ATLAS and CMS have performed searches for pair production of heavy  $B$  quarks which subsequently decay to  $Wt$ ,  $bZ$  or  $bH^0$ . The searches have been carried out in final states with single leptons, di-leptons (with same charge or opposite charge), multileptons, as well as in fully hadronic final states.

##### $B \rightarrow WtX$ :

Search for  $B \rightarrow tW$  has been performed by the ATLAS experiment [116] using lepton+jets events. This analysis relies on a discriminant obtained via the BDT technique. The BDT uses kinematic and topological variables such as the jet and  $b$ -jet multiplicity,  $H_T$ , the angular separation between the lepton and the leading  $b$ -tagged jet or between lepton and hadronic  $W/Z$



**Figure 2:** Top panel: upper limit on the  $T$  quark production cross section for branching fractions into  $bW$ ,  $tH^0$ ,  $tZ$  of 50%, 25%, 25% obtained from the leptonic inclusive analysis [110]. Bottom panel: Branching fraction triangle with observed limits for the  $T$  quark mass from the CMS combination analysis [106].

candidates, the transverse mass of the leptonically decaying  $W$  boson candidate,  $p_T$  of various objects including the leptonically decaying  $W$  boson, the number of hadronic  $W/Z$  candidates, etc. For  $BR(B \rightarrow tW) = 1$ , the lower limit on the mass of the  $B$  quark is obtained to be 810 GeV. For the weak-isospin singlet scenario, a  $B$  quark with mass lower than 640 GeV is excluded. A similar search by CMS [117] selects events with one lepton and four or more jets, with at least one  $b$ -tagged jet, significant missing  $p_T$ , and further categorizes them based on the number of jets tagged as arising from the decay of boosted  $W$ ,  $Z$  or  $H^0$  bosons. The  $S_T$  distributions of the events in different categories show no excess of events above the expected

background and yield a lower limit on the  $B$  quark mass of 732 GeV for  $BR(B \rightarrow Wt) = 1$ .

#### $B \rightarrow bZX$ :

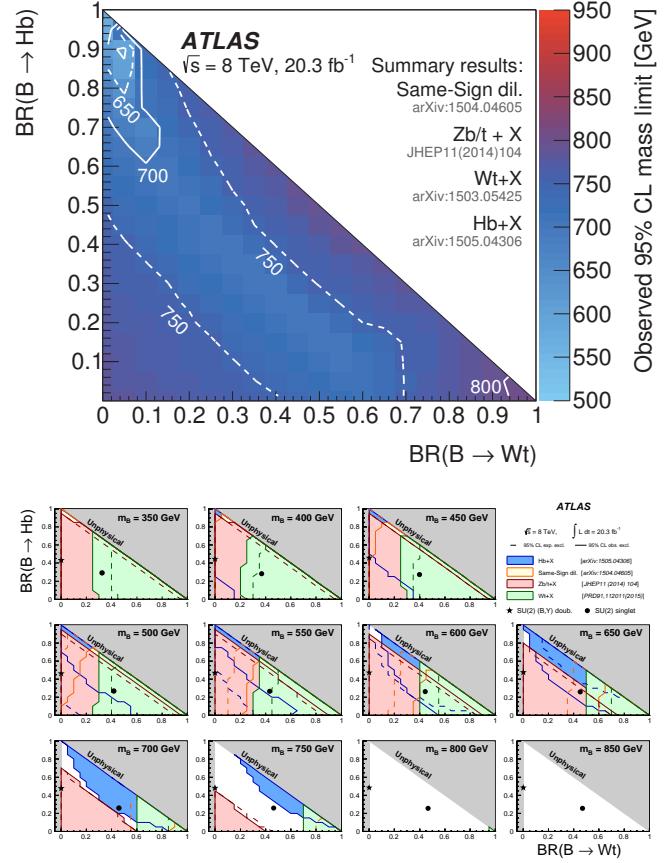
A search by CMS [115] for the pair-production of a heavy  $B$  quark and its antiparticle, one of which decays to  $bZ$  selects events with a  $Z$ -boson decay to  $e^+e^-$  or  $\mu^+\mu^-$  and a jet identified as originating from a  $b$  quark. The signal from  $B \rightarrow bZ$  decays would appear as a local enhancement in the  $bZ$  mass distribution. No such enhancement is found and  $B$  quarks that decay 100% into  $bZ$  are excluded below 700 GeV. This analysis also sets upper limits on the branching fraction for  $B \rightarrow bZ$  decays of 30-100% in the  $B$  quark mass range 450-700 GeV. A complementary search has been carried out by ATLAS for new heavy quarks decaying into a  $Z$  boson and a  $b$ -quark [113]. Selected dilepton events contain a high transverse momentum  $Z$  boson that decays leptonically, together with two  $b$ -jets. If the dilepton events have an extra lepton in addition to those from the  $Z$  boson, then only one  $b$ -jet is required. No significant excess of events above the standard model expectation is observed, and mass limits are set depending on the assumed branching ratios, see Fig. 3. In a weak-isospin singlet scenario, a  $B$  quark with mass lower than 645 GeV is excluded, while for a particular weak-isospin doublet scenario, a  $B$  quark with mass lower than 725 GeV is excluded.

ATLAS has searched for the electroweak production of single  $B$  quarks, which is accompanied by a  $b$ -jet and a light jet [113]. The dilepton selection for double  $B$  production is modified for the single  $B$  production study by requesting the presence of an additional energetic jet in the forward region. An upper limit of 200 fb is obtained for the process  $\sigma(pp \rightarrow B\bar{b}q) \times B(B \rightarrow Zb)$  with a heavy  $B$  quark mass at 700 GeV. This search indicates that the electroweak mixing parameter  $X_{Bb}$  below 0.5 is neither expected or observed to be excluded for any values of  $B$  quark mass.

#### Combination $B \rightarrow tW/bZ/bH^0$ :

The ATLAS experiment has combined the various analyses targeted for specific decay modes to obtain the most sensitive limits on the pair production of  $B$  quarks [109]. The analyses using single lepton events, same sign charge dilepton events, events with opposite sign dilepton events, and multilepton events are combined to obtain lower limits on the mass of the  $B$  quark in the plane of  $BR(B \rightarrow Wt)$  vs  $BR(B \rightarrow bH)$ . The searches are optimized for 100% branching fractions and hence are most sensitive at large  $BR(B \rightarrow Wt)$ , and also at large  $BR(B \rightarrow bH^0)$ . For all possible values of branching ratios in the three decay modes  $tW$ ,  $bZ$ , or  $bH^0$ , the lower limits on the  $B$  quark mass is found to be between 575 GeV and 813 GeV and as shown in Fig. 3.

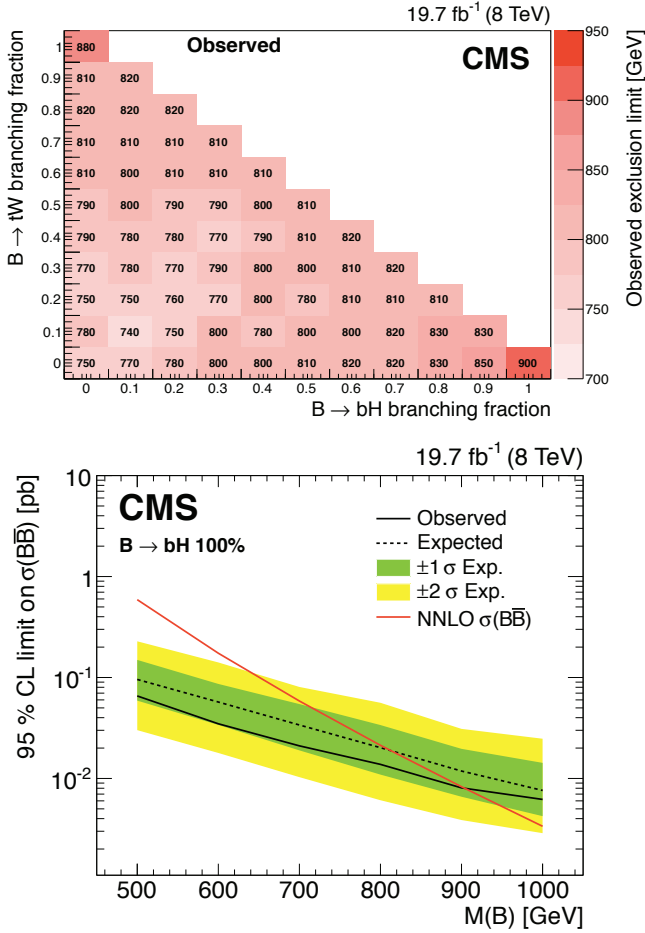
A similar combination of CMS analyses [115] in the final states with single leptons, di-leptons (with same charge or opposite charge), multileptons, as well as fully hadronic decays



**Figure 3:** Observed limits on the mass of the  $T$  quark in the plane of  $BR(B \rightarrow bH^0)$  versus  $BR(B \rightarrow tW)$  from all ATLAS searches for  $BB$  production [109]. Top panel: summary of the most restrictive observed limit on the mass. Contour lines are provided to guide the eye. Bottom panel: Exclusion limits are drawn sequentially for each of the analyses and overlaid (rather than combined).

lead to results shown in Fig. 4. The discriminating variables used in these analyses are  $S_T$ ,  $H_T$  and the invariant mass of the dileptons and the  $b$ -jets. As different topologies target multiple decay modes, with various degree of sensitivity to the  $B$  quark mass, the best results for the  $Wt$  decays is obtained from the combination of lepton+jets, same-sign dilepton and multilepton events, while for the  $bZ$  mode a combination of opposite-sign dilepton and multilepton events leads to the best sensitivity for the mass limits. For the  $bH^0$  decays, the all-hadronic events give the dominant contribution to the mass limit. For  $B$  quarks that decay exclusively into  $tW$  masses below 880 GeV are excluded, while for 100% decay branching fraction of  $B$  to  $bH^0$ ,  $B$  quarks up to 900 GeV are excluded. The exclusion limits for all combinations of branching fractions lie between 740 GeV to 900 GeV, and are shown in Fig. 4, together with the cross section limit plotted for  $B$  quark decays to the  $bH^0$  mode.





**Figure 4:** Top panel: observed limits on the  $B$  quark mass for each combination of branching fractions to  $tW$ ,  $bZ$ , and  $bH^0$  obtained by the combination of channels. The color scale represents the mass exclusion limit obtained at each point [115]. Bottom panel: Observed and expected cross section limit results as a function of  $B$  mass, for the combination of all channels and for exclusive branching fraction of  $B$  to  $bH$  [115].

#### 2.4 A charge +5/3 top-partner quark

In models of dynamical electroweak symmetry breaking, the same interactions which give rise to the mass of the top-quark can give unacceptably large corrections to the branching ratio of the  $Z$  boson to  $b\bar{b}$  [75]. These corrections can be substantially reduced, however, in theories with an extended “custodial symmetry” [45]. This symmetry requires the existence of a charge +5/3 vector-like partner of the top quark.

Both experiments have performed a search for a heavy top vector-like quark  $T_{5/3}$ , with exotic charge 5/3, such as that proposed in Refs. [118,119]. The analyses assume pair-production of  $T_{5/3}$  with  $T_{5/3}$  decaying with 100% branching fraction to  $tW$ . The analysis is based on searching for same-sign leptons, from the two  $W$  bosons from one of the

$T_{5/3}$ . Requiring same-sign leptons eliminates most of the standard model background processes, leaving those with smaller cross sections:  $t\bar{t}$ ,  $W$ ,  $t\bar{t}Z$ ,  $WWW$ , and same-sign  $WW$ . In addition backgrounds from instrumental effects due to charge misidentification are considered. The CMS search also utilizes jet substructure techniques to identify boosted  $T_{5/3}$  topologies. These searches restrict the  $T_{5/3}$  mass to be higher than 800 GeV [120]. The pair-production limits obtained by ATLAS correspond to a lower mass limit on  $T_{5/3}$  of 840 GeV [116].

The single  $T_{5/3}$  production cross section depends on the coupling constant  $\lambda$  of the  $tWT_{5/3}$  vertex. ATLAS has performed an analysis of same-sign dileptons which includes both the single and pair production. This analysis leads to a lower limit on the mass of the  $T_{5/3}$  of 750 GeV for both values of  $\lambda = 0.5$  and  $1.0$  [121].

#### 2.5 Colorons and Colored Scalars

These particles are associated with top-condensate and top-seesaw models, which involve an enlarged color gauge group. The new particles decay to dijets,  $t\bar{t}$ , and  $b\bar{b}$ .

Direct searches for colorons, color-octet scalars and other heavy objects decaying to  $q\bar{q}$ ,  $qg$ ,  $q\bar{q}$ , or  $g\bar{g}$  has been performed using LHC data from pp collisions at  $\sqrt{s} = 7$  and 8 TeV. Based on the analysis of dijet events from a data sample corresponding to a luminosity of  $19.6 \text{ fb}^{-1}$ , the CMS experiment excludes pair production of colorons with mass between  $1.20 - 3.60$  and  $3.90 - 4.08$  TeV at 95% C.L. as shown in Fig. 5 [83]. A search for pair-produced colorons based on an integrated luminosity of  $5.0 \text{ fb}^{-1}$  at  $\sqrt{s} = 7$  TeV by CMS excludes colorons with masses between 250 GeV and 740 GeV, assuming colorons decay 100% into  $q\bar{q}$  [122]. This analysis is based on events with at least four jets and two dijet combinations with similar dijet mass. Color-octet scalars (s8) with masses between  $1.20 - 2.79$  TeV are excluded by CMS (Fig. 5 [83]), and below 2.7 TeV by ATLAS [82].

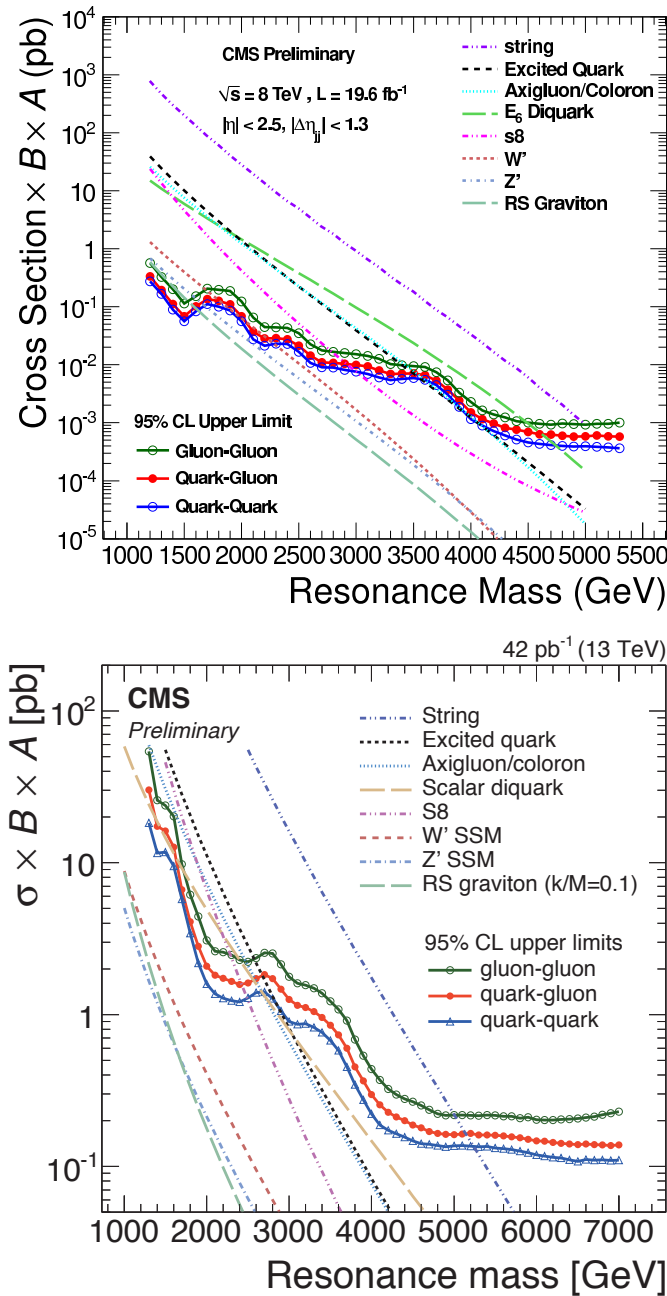
These studies have now been extended to take advantage of the increased center-of-mass energy during Run 2 of the LHC. Using the  $40 \text{ pb}^{-1}$  of data collected at  $\sqrt{s} = 13$  TeV, searches for narrow resonances have been performed by CMS. An analysis of the dijet invariant mass spectrum formed using wide jets [123], separated by  $\Delta\eta_{jj} \leq 1.3$ , leads to limits on new particles decaying to parton pairs ( $q\bar{q}$ ,  $qg$ ,  $g\bar{g}$ ). Specific exclusions on the masses of colorons and color-octet scalars are obtained and shown in Fig. 5.

#### 3. Conclusions

As the above analyses have demonstrated, there is already substantial sensitivity to possible new particles predicted to accompany the  $H^0$  in dynamical frameworks of electroweak symmetry breaking. No hints of any deviations from the standard model have been observed, and limits typically at the scale of a few hundred GeV to 1 TeV are set.

Given the need to better understand the  $H^0$  and to determine in detail how it behaves, we expect that such analyses will be a major theme of Run 2 the LHC, and we look forward to





**Figure 5:** Observed 95% C.L. limits on  $\sigma \times B \times A$  for string resonances, excited quarks, axigluons, colorons, E6 diquarks, s8 resonances,  $W'$  and  $Z'$  bosons, and Randall-Sundrum gravitons  $g_{KK}$ . Top panel: results from Ref. [83] from Run 1. Bottom panel: results from Ref. [123] from Run 2.

increased sensitivity as a result of the higher luminosity at the increased centre of mass energy of collisions.

## References

1. ATLAS Collab., Phys. Lett. **B716**, 1 (2012).
2. CMS Collab., Phys. Lett. **B716**, 30 (2012).

3. S. Weinberg, Physica A **96**, 327 (1979).
4. A. Manohar and H. Georgi, Nucl. Phys. **B234**, 189 (1984).
5. H. Georgi, Nucl. Phys. **B266**, 274 (1986).
6. R.S. Chivukula, hep-ph/0011264 (2000).
7. R.S. Chivukula, M.J. Dugan, and M. Golden, Phys. Rev. **D47**, 2930 (1993).
8. S. Weinberg, Phys. Rev. **D13**, 974 (1976).
9. S. Weinberg, Phys. Rev. **D19**, 1277 (1979).
10. L. Susskind, Phys. Rev. **D20**, 2619 (1979).
11. K. Lane, hep-ph/0202255 (2002).
12. C.T. Hill and E.H. Simmons, Phys. Reports **381**, 235 (2003), [Erratum-ibid. **390**, 553 (2004)].
13. R. Shrock, hep-ph/0703050 (2007).
14. E. Eichten *et al.*, Rev. Mod. Phys. **56**, 579 (1984) [Addendum-ibid. **58**, 1065 (1986)].
15. E. Eichten *et al.*, Phys. Rev. **D34**, 1547 (1986).
16. R.S. Chivukula and V. Koulovassilopoulos, Phys. Lett. **B309**, 371 (1993).
17. R. Foadi, M.T. Frandsen, and F. Sannino, Phys. Rev. **D87**, 095001 (2013).
18. B. Holdom, Phys. Lett. **B150**, 301 (1985).
19. K. Yamawaki, M. Bando, and K.-i. Matumoto, Phys. Rev. Lett. **56**, 1335 (1986).
20. T.W. Appelquist, D. Karabali, and L.C.R. Wijewardhana, Phys. Rev. Lett. **57**, 957 (1986).
21. T. Appelquist and L.C.R. Wijewardhana, Phys. Rev. **D35**, 774 (1987).
22. T. Appelquist and L.C.R. Wijewardhana, Phys. Rev. **D36**, 568 (1987).
23. E. Gildener and S. Weinberg, Phys. Rev. **D13**, 3333 (1976).
24. Z. Chacko, R. Franceschini and R. K. Mishra, JHEP **1304**, 015 (2013).
25. B. Bellazzini *et al.*, Eur. Phys. J. **C73**, 2333 (2013).
26. B. Bellazzini *et al.*, Eur. Phys. J. **C74**, 2790 (2014).
27. Z. Chacko, R.K. Mishra, and D. Stolarski, JHEP **1309**, 121 (2013).
28. J.R. Ellis, M.K. Gaillard, and D.V. Nanopoulos, Nucl. Phys. **B106**, 292 (1976).
29. M.A. Shifman *et al.*, Sov. J. Nucl. Phys. **30**, 711 (1979) [Yad. Fiz. **30**, 1368 (1979)].
30. A.I. Vainshtein, V.I. Zakharov, and M.A. Shifman, Sov. Phys. Usp. **23**, 429 (1980) [Usp. Fiz. Nauk **131**, 537 (1980)].
31. M. Bando, K.-i. Matumoto, and K. Yamawaki, Phys. Lett. **B178**, 308 (1986).
32. W.D. Goldberger, B. Grinstein, and W. Skiba, Phys. Rev. Lett. **100**, 111802 (2008).
33. S. Matsuzaki and K. Yamawaki, Phys. Rev. **D85**, 095020 (2012).
34. S. Matsuzaki and K. Yamawaki, Phys. Rev. **D86**, 035025 (2012).
35. E. Eichten, K. Lane, and A. Martin, arXiv:1210.5462 (2012).
36. D.B. Kaplan and H. Georgi, Phys. Lett. **B136**, 183 (1984).

## Searches Particle Listings

## Technicolor

37. D.B. Kaplan, H. Georgi, and S. Dimopoulos, Phys. Lett. **B136**, 187 (1984).
38. M.E. Peskin, Nucl. Phys. **B175**, 197 (1980).
39. J. Preskill, Nucl. Phys. **B177**, 21 (1981).
40. R. Barbieri and A. Strumia, [hep-ph/0007265](#) (2000).
41. N. Arkani-Hamed, A.G. Cohen, and H. Georgi, Phys. Lett. **B513**, 232 (2001).
42. N. Arkani-Hamed *et al.*, JHEP **0208**, 020 (2002).
43. N. Arkani-Hamed *et al.*, JHEP **0207**, 034 (2002).
44. M. Schmaltz and D. Tucker-Smith, Ann. Rev. Nucl. and Part. Sci. **55**, 229 (2005).
45. K. Agashe *et al.*, Phys. Lett. **B641**, 62 (2006).
46. P. Sikivie *et al.*, Nucl. Phys. **B173**, 189 (1980).
47. R.S. Chivukula, A.G. Cohen, and K.D. Lane, Nucl. Phys. **B343**, 554 (1990).
48. B. Bellazzini, C. Csaki, and J. Serra, Eur. Phys. J. **C74**, 2766 (2014).
49. V.A. Miransky, M. Tanabashi, and K. Yamawaki, Phys. Lett. **B221**, 177 (1989) and Mod. Phys. Lett. **A4**, 1043 (1989).
50. W.A. Bardeen, C.T. Hill, and M. Lindner, Phys. Rev. **D41**, 1647 (1990).
51. C.T. Hill, Phys. Lett. **B266**, 419 (1991).
52. B.A. Dobrescu and C.T. Hill, Phys. Rev. Lett. **81**, 2634 (1998).
53. R.S. Chivukula *et al.*, Phys. Rev. **D59**, 075003 (1999).
54. S. Dimopoulos and L. Susskind, Nucl. Phys. **B155**, 237 (1979).
55. E. Eichten and K.D. Lane, Phys. Lett. **B90**, 125 (1980).
56. D.B. Kaplan, Nucl. Phys. **B365**, 259 (1991).
57. T. Appelquist, M. Piai, and R. Shrock, Phys. Rev. **D69**, 015002 (2004).
58. R.S. Chivukula, B.A. Dobrescu, and E.H. Simmons, Phys. Lett. **B401**, 74 (1997).
59. Y. Grossman and M. Neubert, Phys. Lett. **B474**, 361 (2000).
60. S.J. Huber and Q. Shafi, Phys. Lett. **B498**, 256 (2001).
61. T. Gherghetta and A. Pomarol, Nucl. Phys. **B586**, 141 (2000).
62. K. Agashe, R. Contino, and A. Pomarol, Nucl. Phys. **B719**, 165 (2005).
63. G.F. Giudice *et al.*, JHEP **0706**, 045 (2007).
64. R.S. Chivukula and H. Georgi, Phys. Lett. **B188**, 99 (1987).
65. G. D'Ambrosio *et al.*, Nucl. Phys. **B645**, 155 (2002).
66. K. Agashe *et al.*, [hep-ph/0509117](#) (2005).
67. T. Appelquist and R. Shrock, Phys. Lett. **B548**, 204 (2002).
68. B. Keren-Zur *et al.*, Nucl. Phys. **B867**, 429 (2013).
69. J.M. Maldacena, Adv. Theor. Math. Phys. **2**, 231 (1998).
70. For a review, see C. Csaki, J. Hubisz, and P. Meade, [hep-ph/0510275](#) (2005), and “Extra Dimensions” review in this volume.
71. E.T. Neil, PoS LATTICE **2011**, 009 (2011).
72. J. Giedt, PoS LATTICE **2012**, 006 (2012).
73. J. Kuti, PoS LATTICE **2013**, 004 (2014).
74. T. Appelquist *et al.*, [arXiv:1309.1206 \[hep-lat\]](#).
75. R.S. Chivukula, S.B. Selipsky, and E.H. Simmons, Phys. Rev. Lett. **69**, 575 (1992).
76. ATLAS Collab., Phys. Rev. **D90**, 052005 (2014).
77. CMS Collab., JHEP **1504**, 025 (2015).
78. ATLAS Collab., JHEP **1507**, 157 (2015).
79. CMS Collab., CMS-PAS-EXO-12-046 (2015).
80. ATLAS Collab., CERN-PH-EP-2015-090 (2015), to be published in JHEP.
81. CMS Collab., CMS-B2G-13-008 (2015), submitted to PRD.
82. ATLAS Collab., Phys. Rev. **D91**, 052007 (2015).
83. V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. **D91**, 052009 (2015).
84. ATLAS Collab., ATLAS-CONF-2013-060 (2013).
85. ATLAS Collab., JHEP **1409**, 037 (2014).
86. CMS Collab., CERN-PH-EP-2015-190 (2015).
87. ATLAS Collab., Phys. Lett. **B743**, 235 (2015).
88. ATLAS Collab., Eur. Phys. J. **C75**, 165 (2015).
89. ATLAS Collab., Phys. Lett. **B737**, 223 (2014).
90. CMS Collab., Phys. Lett. **B740**, 83 (2015).
91. ATLAS Collab., Eur. Phys. J. **C75**, 209 (2015).
92. ATLAS Collab., Eur. Phys. J. **C75**, 69 (2015).
93. CMS Collab., JHEP **1408**, 174 (2014).
94. ATLAS Collab., CERN-PH-EP-2015-115 (2015).
95. CMS Collab., JHEP **1408**, 173 (2014).
96. ATLAS Collab., Eur. Phys. J. **C75**, 263 (2015).
97. CMS Collab., CERN-PH-EP-2015-096 (2015), submitted to JHEP.
98. CMS Collab., Phys. Lett. **B748**, 255 (2015).
99. ATLAS Collab., Eur. Phys. J. **C75**, 412 (2015).
100. CMS Collab., CMS-EXO-12-053 (2015).
101. ATLAS Collab., ATLAS-CONF-2013-074 (2013).
102. F. del Aguila *et al.*, Nucl. Phys. **B334**, 1 (1990).
103. CMS Collab., CMS-PAS-B2G-12-017 (2014).
104. S.D. Ellis, C.K. Vermilion, and J.R. Walsh, Phys. Rev. **D80**, 051501 (2009).
105. CMS Collab., CMS-PAS-B2G-12-013 (2015).
106. CMS Collab., submitted to PRD, [arXiv:1509.04177](#).
107. CMS Collab., Phys. Rev. Lett. **107**, 271802 (2011).
108. CMS Collab., Phys. Lett. **B718**, 307 (2012).
109. ATLAS Collab., JHEP **1508**, 105 (2015).
110. CMS Collab., Phys. Lett. **B729**, 149 (2014).
111. CMS Collab., JHEP **1506**, 080 (2015).
112. CMS Collab., [cds.cern.ch/record/1709129](#) (2014).
113. ATLAS Collab., JHEP **1411**, 104 (2014).
114. N. Vignaroli, Phys. Rev. **D86**, 075017 (2012).
115. CMS Collab., submitted to PRD, [arXiv:1507.07129](#).
116. ATLAS Collab., Phys. Rev. **D91**, 112011 (2015).
117. CMS Collab., CMS-PAS-B2G-12-019 (2012), [http://cds.cern.ch/record/1599436](#).
118. R. Contino and G. Servant, JHEP **0806**, 026 (2008).
119. J. Mrazek, A. Wulzer, Phys. Rev. **D81**, 075006 (2010).
120. CMS Collab., Phys. Rev. Lett. **112**, 171801 (2014).
121. ATLAS Collab., JHEP **1510**, 150 (2015).
122. CMS Collab., Phys. Rev. Lett. **110**, 141802 (2013).

123. CMS Collab., CMS-PAS-EXO-15-001 (2015),  
<https://cds.cern.ch/record/2048099>.

The latest unpublished results are described in “Dynamical Electroweak Symmetry Breaking” review.

### MASS LIMITS for Resonances in Models of Dynamical Electroweak Symmetry Breaking

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>2400	95	1 KHACHATRYAN...16E	CMS	top-color $Z'$
		2 AAD 15AB	ATLS	$h \rightarrow \pi_V \pi_V$
>1800	95	3 AAD 15AO	ATLS	top-color $Z'$
		4 AAD 15BB	ATLS	$pp \rightarrow \rho_T / a_{1T} \rightarrow Wh$ or $Zh$
		5 AAD 15Q	ATLS	$h \rightarrow \pi_V \pi_V$
		6 AAIJ 15AN	LHCB	$h \rightarrow \pi_V \pi_V$
>1140	95	7 KHACHATRYAN...15C	CMS	$\rho_T \rightarrow WZ$
		8 KHACHATRYAN...15W	CMS	$H \rightarrow \pi_V \pi_V$
none 200–700, 750–890	95	9 AAD 14AT	ATLS	$pp \rightarrow \omega_T \rightarrow Z\gamma$
none 275–960	95	9 AAD 14AT	ATLS	$pp \rightarrow a_T \rightarrow W\gamma$
		10 AAD 14V	ATLS	color singlet technivector
> 703		11 AAD 13AN	ATLS	$pp \rightarrow a_T \rightarrow W\gamma$
> 494		12 AAD 13AN	ATLS	$pp \rightarrow \omega_T \rightarrow Z\gamma$
none 500–1740	95	13 AAD 13AQ	ATLS	top-color $Z'$
>1300	95	14 CHATRCHYAN 13AP	CMS	top-color $Z'$
>2100	95	13 CHATRCHYAN 13BM	CMS	top-color $Z'$
		15 BAAK 12	RVUE	QCD-like technicolor
none 167–687	95	16 CHATRCHYAN 12AF	CMS	$\rho_T \rightarrow WZ$
> 805	95	13 AALTONEN 11AD	CDF	top-color $Z'$
> 805	95	13 AALTONEN 11AE	CDF	top-color $Z'$
		17 CHIVUKULA 11	RVUE	top-Higgs
		18 CHIVUKULA 11A	RVUE	techni- $\pi$
		19 AALTONEN 10I	CDF	$p\bar{p} \rightarrow \rho_T / \omega_T \rightarrow W\pi_T$
none 208–408	95	20 ABAZOV 10A	D0	$\rho_T \rightarrow WZ$
		21 ABAZOV 07I	D0	$p\bar{p} \rightarrow \rho_T / \omega_T \rightarrow W\pi_T$
> 280	95	22 ABULENCIA 05A	CDF	$\rho_T \rightarrow e^+e^-, \mu^+\mu^-$
		23 CHEKANOV 02B	ZEUS	color octet techni- $\pi$
> 207	95	24 ABAZOV 01B	D0	$\rho_T \rightarrow e^+e^-$
none 90–206.7	95	25 ABDALLAH 01	DLPH	$e^+e^- \rightarrow \rho_T$
		26 AFFOLDER 00F	CDF	color-singlet techni- $\rho$ , $\rho_T \rightarrow W\pi_T, 2\pi_T$
> 600	95	27 AFFOLDER 00K	CDF	color-octet techni- $\rho$ , $\rho_{T8} \rightarrow 2\pi_{LQ}$
none 350–440	95	28 ABE 99F	CDF	color-octet techni- $\rho$ , $\rho_{T8} \rightarrow b\bar{b}$
		29 ABE 99N	CDF	techni- $\omega$ , $\omega_T \rightarrow \gamma b\bar{b}$
none 260–480	95	30 ABE 97G	CDF	color-octet techni- $\rho$ , $\rho_{T8} \rightarrow 2j\text{ets}$

- 1 KHACHATRYAN 16E search for top-color  $Z'$  decaying to  $t\bar{t}$ . The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ . Also exclude  $m_{Z'} < 2.9$  TeV for wider topcolor  $Z'$  with  $\Gamma_{Z'}/m_{Z'} = 0.1$ .
- 2 AAD 15AB search for long-lived hidden valley  $\pi_V$  particles which are produced in pairs by the decay of a scalar boson.  $\pi_V$  is assumed to decay into dijets. See their Fig. 10 for the limit on  $\sigma B$ .
- 3 AAD 15AO search for top-color  $Z'$  decaying to  $t\bar{t}$ . The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ .
- 4 AAD 15BB search for minimal walking technicolor (MWT) isotriplet vector and axial-vector resonances decaying to  $Wh$  or  $Zh$ . See their Fig. 3 for the exclusion limit in the MWT parameter space.
- 5 AAD 15Q search for long-lived hidden valley  $\pi_V$  particles which are produced in pairs by the decay of scalar boson.  $\pi_V$  is assumed to decay into dijets. See their Fig. 5 and Fig. 6 for the limit on  $\sigma B$ .
- 6 AAIJ 15AN search for long-lived hidden valley  $\pi_V$  particles which are produced in pairs by the decay of scalar boson with a mass of 120 GeV.  $\pi_V$  is assumed to decay into dijets. See their Fig. 4 for the limit on  $\sigma B$ .
- 7 KHACHATRYAN 15C search for a vector techni-resonance decaying to  $WZ$ . The limit assumes  $M_{\pi_T} = (3/4) M_{\rho_T} = 25$  GeV. See their Fig. 3 for the limit in  $M_{\pi_T} - M_{\rho_T}$  plane of the low scale technicolor model.
- 8 KHACHATRYAN 15W search for long-lived hidden valley  $\pi_V$  particles which are produced in pairs in the decay of heavy higgs boson  $H$ .  $\pi_V$  is assumed to decay into  $\ell^+\ell^-$ . See their Fig. 7 and Fig. 8 for the limits on  $\sigma B$ .
- 9 AAD 14AT search for techni- $\omega$  and techni- $a$  resonances decaying to  $V\gamma$  with  $V = W(\rightarrow \ell\nu)$  or  $Z(\rightarrow \ell^+\ell^-)$ .
- 10 AAD 14V search for vector techni-resonances decaying into electron or muon pairs in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. See their table IX for exclusion limits with various assumptions.
- 11 AAD 13AN search for vector techni-resonance  $a_T$  decaying into  $W\gamma$ .
- 12 AAD 13AN search for vector techni-resonance  $\omega_T$  decaying into  $Z\gamma$ .
- 13 Search for top-color  $Z'$  decaying to  $t\bar{t}$ . The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ .
- 14 CHATRCHYAN 13AP search for top-color leptophobic  $Z'$  decaying to  $t\bar{t}$ . The quoted limit is for  $\Gamma_{Z'}/m_{Z'} = 0.012$ .
- 15 BAAK 12 give electroweak oblique parameter constraints on the QCD-like technicolor models. See their Fig. 28.
- 16 CHATRCHYAN 12AF search for a vector techni-resonance decaying to  $WZ$ . The limit assumes  $M_{\pi_T} = (3/4) M_{\rho_T} = 25$  GeV. See their Fig. 3 for the limit in  $M_{\pi_T} - M_{\rho_T}$  plane of the low scale technicolor model.

- 17 Using the LHC limit on the Higgs boson production cross section, CHIVUKULA 11 obtain a limit on the top-Higgs mass  $> 300$  GeV at 95% CL assuming 150 GeV top-pion mass.
- 18 Using the LHC limit on the Higgs boson production cross section, CHIVUKULA 11A obtain a limit on the technipion mass ruling out the region  $110 \text{ GeV} < m_P < 2m_t$ . Existence of color techni-fermions, top-color mechanism, and  $N_{TC} \geq 3$  are assumed.
- 19 AALTONEN 10I search for the vector techni-resonances ( $\rho_T, \omega_T$ ) decaying into  $W\pi_T$  with  $W \rightarrow \ell\nu$  and  $\pi_T \rightarrow b\bar{b}, b\bar{c},$  or  $b\bar{u}$ . See their Fig. 3 for the exclusion plot in  $M_{\pi_T} - M_{\rho_T}$  plane.
- 20 ABAZOV 10A search for a vector techni-resonance decaying into  $WZ$ . The limit assumes  $M_{\rho_T} < M_{\pi_T} + M_W$ .
- 21 ABAZOV 07I search for the vector techni-resonances ( $\rho_T, \omega_T$ ) decaying into  $W\pi_T$  with  $W \rightarrow e\nu$  and  $\pi_T \rightarrow b\bar{b}$  or  $b\bar{c}$ . See their Fig. 2 for the exclusion plot in  $M_{\pi_T} - M_{\rho_T}$  plane.
- 22 ABULENCIA 05A search for resonances decaying to electron or muon pairs in  $p\bar{p}$  collisions. at  $\sqrt{s} = 1.96$  TeV. The limit assumes Technicolor-scale mass parameters  $M_V = M_A = 500$  GeV.
- 23 CHEKANOV 02B search for color octet techni- $\pi$   $P$  decaying into dijets in  $ep$  collisions. See their Fig. 5 for the limit on  $\sigma(ep \rightarrow ePX) \cdot B(P \rightarrow 2j)$ .
- 24 ABAZOV 01B searches for vector techni-resonances ( $\rho_T, \omega_T$ ) decaying to  $e^+e^-$ . The limit assumes  $M_{\rho_T} = M_{\omega_T} < M_{\pi_T} + M_W$ .
- 25 The limit is independent of the  $\pi_T$  mass. See their Fig. 9 and Fig. 10 for the exclusion plot in the  $M_{\rho_T} - M_{\pi_T}$  plane. ABDALLAH 01 limit on the technipion mass is  $M_{\pi_T} > 79.8$  GeV for  $N_D = 2$ , assuming its point-like coupling to gauge bosons.
- 26 AFFOLDER 00F search for  $\rho_T$  decaying into  $W\pi_T$  or  $\pi_T\pi_T$  with  $W \rightarrow \ell\nu$  and  $\pi_T \rightarrow b\bar{b}, b\bar{c}$ . See Fig. 1 in the above Note on “Dynamical Electroweak Symmetry Breaking” for the exclusion plot in the  $M_{\rho_T} - M_{\pi_T}$  plane.
- 27 AFFOLDER 00K search for the  $\rho_{T8}$  decaying into  $\pi_{LQ}\pi_{LQ}$  with  $\pi_{LQ} \rightarrow b\nu$ . For  $\pi_{LQ} \rightarrow c\nu$ , the limit is  $M_{\rho_{T8}} > 510$  GeV. See their Fig. 2 and Fig. 3 for the exclusion plot in the  $M_{\rho_{T8}} - M_{\pi_{LQ}}$  plane.
- 28 ABE 99F search for a new particle  $X$  decaying into  $b\bar{b}$  in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. See Fig. 7 in the above Note on “Dynamical Electroweak Symmetry Breaking” for the upper limit on  $\sigma(p\bar{p} \rightarrow X) \times B(X \rightarrow b\bar{b})$ . ABE 99F also exclude top gluons of width  $\Gamma = 0.3M$  in the mass interval  $280 < M < 670$  GeV, of width  $\Gamma = 0.5M$  in the mass interval  $340 < M < 640$  GeV, and of width  $\Gamma = 0.7M$  in the mass interval  $375 < M < 560$  GeV.
- 29 ABE 99N search for the techni- $\omega$  decaying into  $\gamma\pi_T$ . The technipion is assumed to decay  $\pi_T \rightarrow b\bar{b}$ . See Fig. 2 in the above Note on “Dynamical Electroweak Symmetry Breaking” for the exclusion plot in the  $M_{\omega_T} - M_{\pi_T}$  plane.
- 30 ABE 97G search for a new particle  $X$  decaying into dijets in  $p\bar{p}$  collisions at  $E_{cm} = 1.8$  TeV. See Fig. 5 in the above Note on “Dynamical Electroweak Symmetry Breaking” for the upper limit on  $\sigma(p\bar{p} \rightarrow X) \times B(X \rightarrow 2j)$ .

### REFERENCES FOR Technicolor

KHACHATRYAN... 16E	PR D93 012001	V. Khachatryan et al.	(CMS Collab.)
AAD 15AB	PR D92 012010	G. Aad et al.	(ATLAS Collab.)
AAD 15AO	JHEP 1508 148	G. Aad et al.	(ATLAS Collab.)
AAD 15BB	EPJ C75 263	G. Aad et al.	(ATLAS Collab.)
AAD 15Q	PL B743 15	G. Aad et al.	(ATLAS Collab.)
AAIJ 15AN	EPJ C75 152	R. Aaij et al.	(LHCb Collab.)
KHACHATRYAN... 15C	PL B740 83	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN... 15W	PR D91 052012	V. Khachatryan et al.	(CMS Collab.)
AAD 14AT	PL B738 428	G. Aad et al.	(ATLAS Collab.)
AAD 14V	PR D90 052005	G. Aad et al.	(ATLAS Collab.)
AAD 13AN	PR D87 112003	G. Aad et al.	(ATLAS Collab.)
Also	PR D91 119901 (err.)	G. Aad et al.	(ATLAS Collab.)
AAD 13AQ	PR D88 012004	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN 13AP	PR D87 072002	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13BM	PRL 111 211804	S. Chatrchyan et al.	(CMS Collab.)
Also	PRL 112 119903 (err.)	S. Chatrchyan et al.	(CMS Collab.)
BAAK 12	EPJ C72 2003	M. Baak et al.	(Glitter Group)
CHATRCHYAN 12AF	PRL 109 141801	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN 11AD	PR D84 072003	T. Aaltonen et al.	(CDF Collab.)
AALTONEN 11AE	PR D84 072004	T. Aaltonen et al.	(CDF Collab.)
CHIVUKULA 11	PR D84 095022	R.S. Chivukula et al.	(CDF Collab.)
CHIVUKULA 11A	PR D84 115025	R. S. Chivukula et al.	(CDF Collab.)
AALTONEN 10I	PRL 104 111802	T. Aaltonen et al.	(CDF Collab.)
ABAZOV 10A	PRL 104 061801	V.M. Abazov et al.	(DO Collab.)
ABAZOV 07I	PRL 98 221801	V.M. Abazov et al.	(DO Collab.)
ABULENCIA 05A	PRL 95 252001	A. Abulencia et al.	(CDF Collab.)
CHEKANOV 02B	PL B531 9	S. Chekanov et al.	(ZEUS Collab.)
ABAZOV 01B	PRL 87 061802	V.M. Abazov et al.	(DO Collab.)
ABDALLAH 01	EPJ C22 17	J. Abdallah et al.	(DELPHI Collab.)
AFFOLDER 00F	PRL 84 1110	T. Affolder et al.	(CDF Collab.)
AFFOLDER 00K	PRL 85 2056	T. Affolder et al.	(CDF Collab.)
ABE 99F	PRL 82 2038	F. Abe et al.	(CDF Collab.)
ABE 99N	PRL 83 3124	F. Abe et al.	(CDF Collab.)
ABE 97G	PR D55 R5263	F. Abe et al.	(CDF Collab.)

# Searches Particle Listings

## Quark and Lepton Compositeness

### Quark and Lepton Compositeness, Searches for

The latest unpublished results are described in the “Quark and Lepton Compositeness” review.

#### SEARCHES FOR QUARK AND LEPTON COMPOSITENESS

Revised 2015 by K. Hikasa (Tohoku University), M. Tanabashi (Nagoya University), K. Terashi (University of Tokyo), and N. Varelas (University of Illinois at Chicago)

#### Limits on contact interactions

If quarks and leptons are made of constituents, then at the scale of constituent binding energies there should appear new interactions among them. At energies much below the compositeness scale ( $\Lambda$ ), these interactions are suppressed by inverse powers of  $\Lambda$ . The dominant effect of the compositeness of fermion  $\psi$  should come from the lowest dimensional interactions with four fermions (contact terms), whose most general flavor-diagonal color-singlet chirally invariant form reads [1,2]

$$\mathcal{L} = \frac{g_{\text{contact}}^2}{2\Lambda^2} \sum_{i,j} \left[ \eta_{LL}^{ij} (\bar{\psi}_L^i \gamma_\mu \psi_L^i) (\bar{\psi}_L^j \gamma^\mu \psi_L^j) + \eta_{RR}^{ij} (\bar{\psi}_R^i \gamma_\mu \psi_R^i) (\bar{\psi}_R^j \gamma^\mu \psi_R^j) + \eta_{LR}^{ij} (\bar{\psi}_L^i \gamma_\mu \psi_L^i) (\bar{\psi}_R^j \gamma^\mu \psi_R^j) + \eta_{RL}^{ij} (\bar{\psi}_R^i \gamma_\mu \psi_R^i) (\bar{\psi}_L^j \gamma^\mu \psi_L^j) \right], \quad (1)$$

with  $i, j$  being the indices of fermion species. Color and other indices are suppressed in Eq. (1). Chiral invariance provides a natural explanation why quark and lepton masses are much smaller than their inverse size  $\Lambda$ . Note  $\eta_{\alpha\beta}^{ij} = \eta_{\beta\alpha}^{ji}$ , therefore, in order to specify the contact interaction among the same fermion species  $i = j$ , it is enough to use  $\eta_{LL}$ ,  $\eta_{RR}$  and  $\eta_{LR}$ . We will suppress the indices of fermion species hereafter. We may determine the scale  $\Lambda$  unambiguously by using the above form of the effective interactions; the conventional method [1] is to fix its scale by setting  $g_{\text{contact}}^2/4\pi = g_{\text{contact}}^2(\Lambda)/4\pi = 1$  for the new strong interaction coupling and by setting the largest magnitude of the coefficients  $\eta_{\alpha\beta}$  to be unity. In the following, we denote

$$\begin{aligned} \Lambda &= \Lambda_{LL}^\pm \text{ for } (\eta_{LL}, \eta_{RR}, \eta_{LR}) = (\pm 1, 0, 0), \\ \Lambda &= \Lambda_{RR}^\pm \text{ for } (\eta_{LL}, \eta_{RR}, \eta_{LR}) = (0, \pm 1, 0), \\ \Lambda &= \Lambda_{VV}^\pm \text{ for } (\eta_{LL}, \eta_{RR}, \eta_{LR}) = (\pm 1, \pm 1, \pm 1), \\ \Lambda &= \Lambda_{AA}^\pm \text{ for } (\eta_{LL}, \eta_{RR}, \eta_{LR}) = (\pm 1, \pm 1, \mp 1), \\ \Lambda &= \Lambda_{V-A}^\pm \text{ for } (\eta_{LL}, \eta_{RR}, \eta_{LR}) = (0, 0, \pm 1). \end{aligned} \quad (2)$$

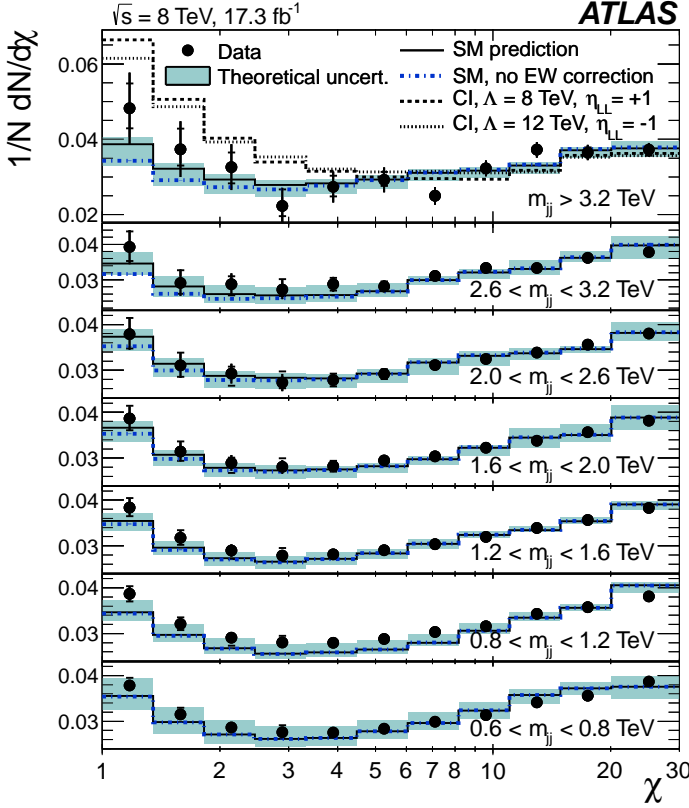
Such interactions can arise by interchanging constituents (when the fermions have common constituents), and/or by exchanging the binding quanta (whenever binding quanta couple to constituents of both particles).

Fermion scattering amplitude induced from the contact interaction in Eq. (1) interferes with the Standard Model (SM)

amplitude destructively or constructively. The sign of interference depends on the sign of  $\eta_{\alpha\beta}$ . For instance, in the parton level  $qq \rightarrow qq$  scattering cross section in the  $\Lambda_{LL}^\pm$  model, the contact interaction amplitude and the SM gluon exchange amplitude interfere destructively for  $\eta_{LL} = +1$ , while they interfere constructively for  $\eta_{LL} = -1$ . In models of quark compositeness, the quark scattering cross sections induced from the contact interactions receive sizable QCD radiative corrections. Ref. 3 provides the exact next-to-leading order (NLO) QCD corrections to the contact interaction induced quark scattering cross sections.

Over the last three decades experiments at the CERN SpS [4,5], the Fermilab Tevatron [6,7], and the CERN LHC [8–12] have searched for quark contact interactions, characterized by the four-fermion effective Lagrangian in Eq. (1), using jet final states. These searches have been performed primarily by studying the angular distribution of the two highest transverse momentum,  $p_T$ , jets (dijets), and the inclusive jet  $p_T$  spectrum. The variable  $\chi = \exp(|(y_1 - y_2)|)$  is used to measure the dijet angular distribution, where  $y_1$  and  $y_2$  are the rapidities of the two jets with the highest transverse momenta. For collinear massless parton scattering,  $\chi$  is related to the polar scattering angle  $\theta^*$  in the partonic center-of-mass frame by  $\chi = (1 + |\cos \theta^*|)/(1 - |\cos \theta^*|)$ . The choice of  $\chi$  is motivated by the fact that the angular distribution for Rutherford scattering, which is proportional to  $1/(1 - \cos \theta^*)^2$ , is independent of  $\chi$ . In perturbative QCD the  $\chi$  distributions are relatively flat and only mildly modified by higher-order QCD or electroweak corrections. Signatures of quark contact interactions exhibit more isotropic angular distribution than QCD and they can be identified as an excess at low values of  $\chi$ . In the inclusive jet cross section measurement, quark contact interaction effects are searched as deviations from the predictions of perturbative QCD in the tails of the high- $p_T$  jet spectrum.

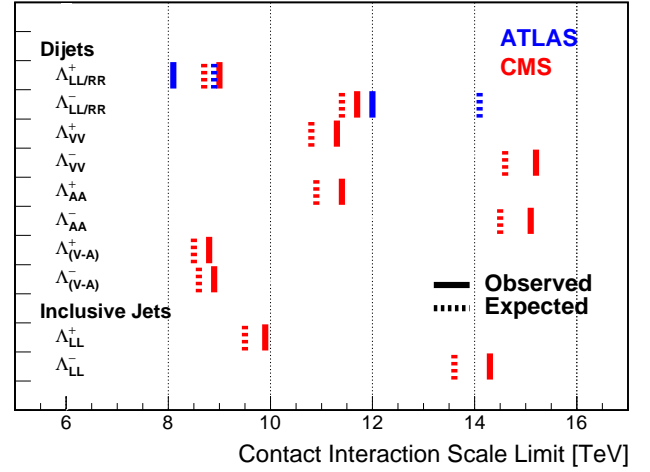
Recent results from the LHC, using data collected at proton-proton center-of-mass energies of  $\sqrt{s} = 7$  and 8 TeV, extend previous Tevatron limits on quark contact interactions. Figure 1 shows the normalized dijet angular distributions for several dijet mass ranges measured in ATLAS [9] at  $\sqrt{s} = 8$  TeV. The data distributions are compared with SM predictions, estimated using PYTHIA8 [13] with GEANT4-based [14] ATLAS detector simulation and corrected to NLO QCD calculation provided by NLO Jet++ [15] including electroweak corrections [16], and with predictions including a contact interaction term in which only left-handed quarks participate at compositeness scale  $\Lambda_{LL}^+ = 8$  TeV ( $\Lambda_{LL}^- = 12$  TeV) with destructive (constructive) interference. Over a wide range of  $\chi$  and dijet mass the data are well described by the SM predictions. Using the dijet angular distributions measured at high dijet masses and  $\sqrt{s} = 8$  TeV, the ATLAS [9] and CMS [12] Collaborations have set 95% confidence level (C.L.) lower limits on the contact interaction scale  $\Lambda$ , ranging from 8.1 to 15.2 TeV for different quark contact interaction models that correspond to various combinations of  $(\eta_{LL}, \eta_{RR}, \eta_{LR})$ , as summarized in Figure 2.



**Figure 1:** Normalized dijet angular distributions in several dijet mass ( $m_{jj}$ ) ranges. The data distributions are compared to the SM predictions (solid line) and with the predictions including a contact interaction (CI) term in which only left-handed quarks participate of compositeness scale  $\Lambda_{LL}^+ = 8$  TeV (dashed line) and  $\Lambda_{LL}^- = 12$  TeV (dotted line). The SM prediction without the electroweak (EW) corrections is also shown (blue dashed dotted line). The error bars on the data points represent statistical and experimental uncertainties combined in quadrature. The ticks on the error bars represent experimental uncertainties only. The shaded band displayed around the SM prediction shows the theoretical uncertainties. Figure adopted from Ref. 9.

The contact interaction scale limits extracted using the dijet angular distributions include the exact NLO QCD corrections to dijet production induced by contact interactions [3]. In proton-proton collisions, the  $\Lambda_{LL}^\pm$  and  $\Lambda_{RR}^\pm$  contact interaction models result in identical tree-level cross sections and NLO QCD corrections and yield the same exclusion limits. For  $\Lambda_{VV}^\pm$  and  $\Lambda_{AA}^\pm$ , the contact interaction predictions are identical at tree level, but exhibit different NLO QCD corrections and yield different exclusion limits. Figure 2 also shows lower limits for two benchmark contact interaction models in which only left-handed quarks participate with destructive ( $\eta_{LL} = +1$ ) and constructive ( $\eta_{LL} = -1$ ) interference, using the inclusive jet  $p_T$  spectrum measured in CMS at  $\sqrt{s} = 7$  TeV [11].

If leptons ( $l$ ) and quarks ( $q$ ) are composite with common constituents, the interaction of these constituents will manifest itself in the form of a  $llqq$ -type four-fermion contact interaction Lagrangian at energies below the compositeness scale  $\Lambda$ . The  $llqq$  terms in the contact interaction Lagrangian can be expressed as



**Figure 2:** Observed (solid lines) and expected (dashed lines) 95% C.L. lower limits on the contact interaction scale  $\Lambda$  for different contact interaction models from ATLAS [9] and CMS [11,12] using the dijet angular distributions and the inclusive jet  $p_T$  spectrum. The contact interaction models used for the dijet angular distributions include the exact NLO QCD corrections to dijet production. All limits are extracted using the  $CL_s$  technique [17,18].

$$\mathcal{L} = \frac{g_{\text{contact}}^2}{\Lambda^2} [\eta_{LL}(\bar{q}_L \gamma_\mu q_L)(\bar{l}_L \gamma^\mu l_L) + \eta_{RR}(\bar{q}_R \gamma_\mu q_R)(\bar{l}_R \gamma^\mu l_R) + \eta_{LR}(\bar{q}_L \gamma_\mu q_L)(\bar{l}_R \gamma^\mu l_R) + \eta_{RL}(\bar{q}_R \gamma_\mu q_R)(\bar{l}_L \gamma^\mu l_L)]. \quad (3)$$

Searches on quark-lepton compositeness have been reported from experiments at LEP [19–23], HERA [24,25], the Tevatron [26–30], and recently from the ATLAS [31–34] and CMS [35] experiments at the LHC. The most stringent searches for  $llqq$  contact interactions are performed by the LHC experiments using high-mass oppositely-charged lepton pairs produced through the  $q\bar{q} \rightarrow l^+l^-$  Drell-Yan process. The contact interaction amplitude of the  $u\bar{u} \rightarrow l^+l^-$  process interferes with the corresponding SM amplitude constructively (destructively) for  $\eta_{\alpha\beta} = -1$  ( $\eta_{\alpha\beta} = +1$ ). The ATLAS Collaboration has extracted limits on the  $llqq$  contact interaction for the right-right ( $\eta_{RR} = \pm 1$ ,  $\eta_{LL} = \eta_{LR} = \eta_{RL} = 0$ ), left-left ( $\eta_{LL} = \pm 1$ ,  $\eta_{RR} = \eta_{LR} = \eta_{RL} = 0$ ), and left-right ( $\eta_{LR} = \eta_{RL} = \pm 1$ ,  $\eta_{RR} = \eta_{LL} = 0$ ) models [34]. With the ATLAS full dataset at  $\sqrt{s} = 8$  TeV and combining the dielectron and dimuon channels, the 95% C.L. lower limits on the  $llqq$  contact interaction scale  $\Lambda$  are 21.1 TeV (17.5 TeV) for the right-right model, 21.6 TeV (17.2 TeV) for the left-left model, and 26.3

# Searches Particle Listings

## Quark and Lepton Compositeness

TeV (19.0 TeV) for the left-right model, each with constructive (destructive) interference [34]. The limits are extracted using a Bayesian approach with a prior probability flat in  $1/\Lambda^2$ . Using the dimuon channel from the 7-TeV run, the CMS Collaboration, using the  $CL_s$  technique, has set a 95% C.L. lower limit on the scale  $\Lambda$  of 13.1 TeV (9.5 TeV) for the benchmark left-left  $llqq$  contact interaction model with constructive (destructive) interference [35].

Note that the contact interactions arising from the compositeness of quarks and leptons Eq. (1) can also be regarded as a part of more general dimension six operators in the context of low energy standard model effective theory. For a complete list of these dimension six operators see Refs. 36,37.

Interactions of hypothetical dark matter candidate particles with SM can also be described as contact interactions at low energy. See “*Searches for WIMPs and Other Particles*” in this volume for limits on the interactions involving dark matter candidate particles.

### Limits on excited fermions

Another typical consequence of compositeness is the appearance of excited leptons and quarks ( $l^*$  and  $q^*$ ). Phenomenologically, an excited lepton is defined to be a heavy lepton which shares a leptonic quantum number with one of the existing leptons (an excited quark is defined similarly). For example, an excited electron  $e^*$  is characterized by a nonzero transition-magnetic coupling with electrons. Smallness of the lepton mass and the success of QED prediction for  $g - 2$  suggest chirality conservation, *i.e.*, an excited lepton should not couple to both left- and right-handed components of the corresponding lepton [38–40].

Excited leptons may be classified by  $SU(2) \times U(1)$  quantum numbers. Typical examples are:

1. Sequential type  

$$\begin{pmatrix} \nu^* \\ l^* \end{pmatrix}_L, \quad [\nu_R^*], \quad l_R^*.$$

$\nu_R^*$  is necessary unless  $\nu^*$  has a Majorana mass.

2. Mirror type  

$$[\nu_L^*], \quad l_L^*, \quad \begin{pmatrix} \nu^* \\ l^* \end{pmatrix}_R.$$

3. Homodoublet type  

$$\begin{pmatrix} \nu^* \\ l^* \end{pmatrix}_L, \quad \begin{pmatrix} \nu^* \\ l^* \end{pmatrix}_R.$$

Similar classification can be made for excited quarks.

Excited fermions can be pair produced via their minimal gauge couplings. The couplings of excited leptons with  $Z$  are given by

$$\begin{aligned} & \frac{e}{2 \sin \theta_W \cos \theta_W} (-1 + 2 \sin^2 \theta_W) \bar{l}^* \gamma^\mu l^* Z_\mu \\ & + \frac{e}{2 \sin \theta_W \cos \theta_W} \bar{\nu}^* \gamma^\mu \nu^* Z_\mu \end{aligned}$$

in the homodoublet model. The corresponding couplings of excited quarks can be easily obtained. Although form factor effects can be present for the gauge couplings at  $q^2 \neq 0$ , they are usually neglected.

Excited fermions may also be produced via the contact interactions with ordinary quarks and leptons [41]

$$\begin{aligned} \mathcal{L} = & \frac{g_{\text{contact}}^2}{\Lambda^2} [\eta'_{LL} (\bar{\psi}_L \gamma_\mu \psi_L) (\bar{\psi}_L^* \gamma^\mu \psi_L^*) \\ & + (\eta''_{LL} (\bar{\psi}_L \gamma_\mu \psi_L) (\bar{\psi}_L^* \gamma^\mu \psi_L) + \text{h.c.}) + \dots] \end{aligned} \quad (4)$$

Again, the coefficient is conventionally taken  $g_{\text{contact}}^2 = 4\pi$ . It is widely assumed  $\eta'_{LL} = \eta''_{LL} = 1$ ,  $\eta'_{LR} = \eta''_{LR} = \eta'_{RL} = \eta''_{RL} = \eta'_{RR} = \eta''_{RR} = 0$  in experimental analyses for simplicity.

In addition, transition-magnetic type couplings with a gauge boson are expected. These couplings can be generally parameterized as follows:

$$\begin{aligned} \mathcal{L} = & \frac{\lambda_\gamma^{(\psi^*)} e}{2m_{\psi^*}} \bar{\psi}^* \sigma^{\mu\nu} (\eta_L \frac{1-\gamma_5}{2} + \eta_R \frac{1+\gamma_5}{2}) \psi F_{\mu\nu} \\ & + \frac{\lambda_Z^{(\psi^*)} e}{2m_{\psi^*}} \bar{\psi}^* \sigma^{\mu\nu} (\eta_L \frac{1-\gamma_5}{2} + \eta_R \frac{1+\gamma_5}{2}) \psi Z_{\mu\nu} \\ & + \frac{\lambda_W^{(l^*)} g}{2m_{l^*}} \bar{l}^* \sigma^{\mu\nu} \frac{1-\gamma_5}{2} \nu W_{\mu\nu} \\ & + \frac{\lambda_W^{(\nu^*)} g}{2m_{\nu^*}} \bar{\nu}^* \sigma^{\mu\nu} (\eta_L \frac{1-\gamma_5}{2} + \eta_R \frac{1+\gamma_5}{2}) l W_{\mu\nu}^\dagger \\ & + \text{h.c.}, \end{aligned} \quad (5)$$

where  $g = e/\sin \theta_W$ ,  $\psi = \nu$  or  $l$ ,  $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$  is the photon field strength,  $Z_{\mu\nu} = \partial_\mu Z_\nu - \partial_\nu Z_\mu$ , *etc.*. The normalization of the coupling is chosen such that

$$\max(|\eta_L|, |\eta_R|) = 1.$$

Chirality conservation requires

$$\eta_L \eta_R = 0. \quad (6)$$

These couplings in Eq. (5) can arise from  $SU(2) \times U(1)$ -invariant higher-dimensional interactions. A well-studied model is the interaction of homodoublet type  $l^*$  with the Lagrangian (see Refs. 42,43)

$$\mathcal{L} = \frac{1}{2\Lambda} \bar{L}^* \sigma^{\mu\nu} (g f \frac{\tau^a}{2} W_{\mu\nu}^a + g' f' Y B_{\mu\nu}) \frac{1-\gamma_5}{2} L + \text{h.c.}, \quad (7)$$

where  $L$  denotes the lepton doublet  $(\nu, l)$ ,  $\Lambda$  is the compositeness scale,  $g, g'$  are  $SU(2)$  and  $U(1)_Y$  gauge couplings, and  $W_{\mu\nu}^a$  and  $B_{\mu\nu}$  are the field strengths for  $SU(2)$  and  $U(1)_Y$  gauge fields. These couplings satisfy the relation

$$\lambda_W = -\sqrt{2} \sin^2 \theta_W (\lambda_Z \cot \theta_W + \lambda_\gamma), \quad (8)$$

with  $\lambda_{W,Z,\gamma}$  being defined in Eq. (5) with  $\lambda_{W,Z,\gamma} = \lambda_{W,Z,\gamma}^{(\ell^*)}$  or  $\lambda_{W,Z,\gamma} = \lambda_{W,Z,\gamma}^{(\nu^*)}$ . Here  $(\eta_L, \eta_R) = (1, 0)$  is assumed. It should be noted that the electromagnetic radiative decay of  $l^*$  ( $\nu^*$ ) is forbidden if  $f = -f'$  ( $f = f'$ ).

Additional coupling with gluons is possible for excited quarks:

$$\begin{aligned} \mathcal{L} = & \frac{1}{2\Lambda} \bar{Q}^* \sigma^{\mu\nu} \left( g_s f_s \frac{\lambda^a}{2} G_{\mu\nu}^a + g f \frac{\tau^a}{2} W_{\mu\nu}^a + g' f' Y B_{\mu\nu} \right) \\ & \times \frac{1-\gamma_5}{2} Q + \text{h.c.}, \end{aligned} \quad (9)$$

where  $Q$  denotes a quark doublet,  $g_s$  is the QCD gauge coupling, and  $G_{\mu\nu}^a$  the gluon field strength.

If leptons are made of color triplet and antitriplet constituents, we may expect their color-octet partners. Transitions between the octet leptons ( $l_8$ ) and the ordinary lepton ( $l$ ) may take place via the dimension-five interactions

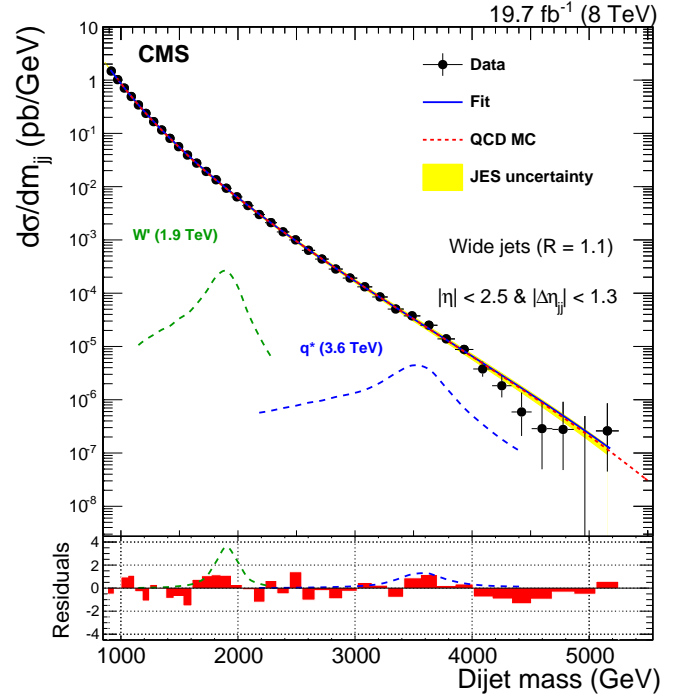
$$\mathcal{L} = \frac{1}{2\Lambda} \sum_l \{ \bar{l}_8^\alpha g_S F_{\mu\nu}^{\alpha} \sigma^{\mu\nu} (\eta_L l_L + \eta_R l_R) + \text{h.c.} \} \quad (10)$$

where the summation is over charged leptons and neutrinos. The leptonic chiral invariance implies  $\eta_L \eta_R = 0$  as before.

Searches for excited quarks and leptons have been performed over the last decades in experiments at the LEP [44–52], HERA [53–56], Tevatron [57–62], and LHC [63–80]. Most stringent constraints from these experiments described below are all given at 95% confidence level.

The signature of excited quarks  $q^*$  at hadron colliders is characterized by a narrow resonant peak in the reconstructed invariant mass distribution of  $q^*$  decay products. The decays via the transition-magnetic type operator in Eq. (9) are considered for excited quarks in LHC searches, and the final states to search for are dijet ( $qg$ ) [63–65, 71–74] or a jet in association with a photon ( $q\gamma$ ) [66, 67, 75] or a weak gauge boson ( $qW, qZ$ ) [76–78]. All analyses consider only spin-1/2 excited states of first generation quarks ( $u^*, d^*$ ) with degenerate masses, expected to be predominantly produced in proton-proton collisions, except for Ref. 74 where excited  $b$  quarks are also considered. Only the minimal gauge interactions and the transition-magnetic couplings with the form given in Eq. (9) are considered in the production process, and hence the contact interactions in Eq. (4) are not considered. The compositeness scale  $\Lambda$  is taken to be the same as the excited quark mass  $m_{q^*}$ . The transition-magnetic coupling coefficients  $f_s, f$  and  $f'$  are assumed to be equal (denoted by  $f$ ) and around order 1.

With the full proton-proton collision data recorded at  $\sqrt{s} = 8$  TeV at LHC, the excited quark masses are excluded in dijet resonance searches up to 4.06 TeV in ATLAS [65] and 3.5 TeV in CMS [74]. Figure 3 shows the dijet mass distribution measured in CMS by using the two highest  $p_T$  jets reconstructed with the anti- $k_T$  algorithm [81] of a distance parameter of 0.5, and by combining nearby jets within  $\Delta R = \sqrt{\Delta\eta^2 + \Delta\phi^2} < 1.1$  around the leading two jets. The measured dijet mass spectrum is compared to a fit with smoothly falling background shape (solid curve) to look for a narrow resonance (3.6 TeV excited quark signal shown as one of two benchmark signals) and predictions from multi-jet events (dashed curve labeled as QCD MC) generated using PYTHIA 6.426 [82] with GEANT4-based [14] CMS detector simulation. The photon + jet resonance searches have excluded excited quarks with mass up to 3.5 TeV in both ATLAS [67] and CMS [75]. All these mass exclusions are obtained for  $f = 1$ . The  $W/Z$  boson + jet final states are examined to look for  $q^* \rightarrow q + W$  and  $q + Z$  signal in CMS [78], exploiting jet substructure technique designed to provide sensitivity for highly-boosted hadronically decaying  $W$



**Figure 3:** Dijet mass distribution measured by CMS using wide jets reconstructed from two highest transverse momentum jets by adding nearby jets within  $\Delta R = \sqrt{\Delta\eta^2 + \Delta\phi^2} < 1.1$ . The data distribution is compared to a fit representing a smooth background spectrum (solid curve) and to the normalized prediction of multi-jet background simulated by PYTHIA (labeled as QCD MC). Excited quark signal with mass of 3.6 TeV is also shown for comparison. Shown at the bottom panel is the bin-by-bin fit residuals normalized by the statistical uncertainty of the data. Figure adopted from Ref. 74.

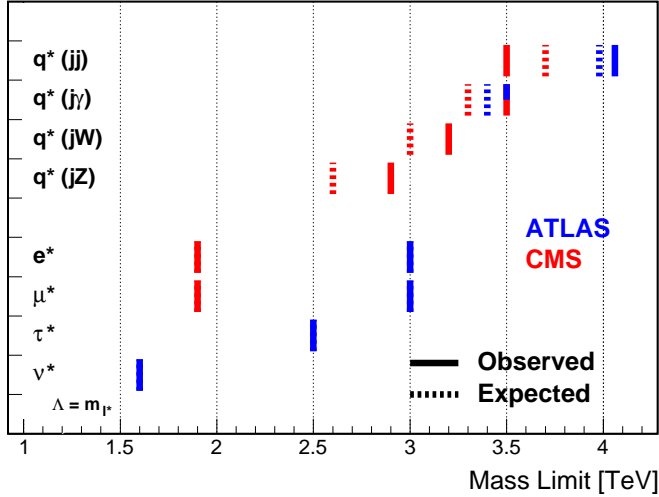
and  $Z$  bosons. The  $q^*$  mass exclusion of 3.2 (2.9) TeV is obtained from the  $W + \text{jet}$  ( $Z + \text{jet}$ ) search.

Searches for excited leptons  $l^*$  are also performed at the LHC using proton-proton collision data recorded at  $\sqrt{s} = 7$  and 8 TeV [68–70, 79, 80]. Considering the single  $l^*$  production Eq. (4) and electromagnetic radiative decay to a SM lepton ( $l$ ) and a photon ( $\gamma$ ), both the excited electron and excited muon masses below 2.2 TeV are excluded for  $\Lambda = m_{l^*}$  at  $\sqrt{s} = 8$  TeV in ATLAS [69]. With the full data at  $\sqrt{s} = 8$  TeV, the inclusive search on multi-lepton signatures with 3 or more charged leptons in ATLAS [70] further constrains the excited charged leptons and neutrinos. Considering both the transition-magnetic Eq. (7) and contact interaction Eq. (4) processes, the lower mass limits for the  $e^*, \mu^*, \tau^*$  and  $\nu^*$  (for every excited neutrino flavor) are obtained to be 3.0, 3.0, 2.5 and 1.6 TeV, respectively, for  $\Lambda = m_{e^*}, m_{\mu^*}, m_{\tau^*}$  and  $m_{\nu^*}$ . The rate of pair-produced excited leptons is independent of  $\Lambda$  for the minimal gauge interaction processes, and it allows to improve the search sensitivity with multi-lepton signatures at high  $\Lambda$ , especially for excited neutrinos because the predominant  $\nu_l^* \rightarrow l + W$  decays



# Searches Particle Listings

## Quark and Lepton Compositeness



**Figure 4:** 95% C.L. lower mass limits for the excited quarks and leptons at ATLAS [65,67,69,70] and CMS [74,75,78,80] experiments. Shown are the most stringent limits for each excited fermion from both experiments. Only first generation quarks ( $u, d$ ) with transition-magnetic type interactions with  $f_s = f = f' = 1$  are considered for excited quarks, and the limits are shown for different final states denoted in parentheses. Excited lepton limits are given for contact interactions with  $\Lambda = m_{l^*}$ . The observed limit from  $q^* \rightarrow q + \gamma$  is 3.5 TeV for both ATLAS and CMS. For the excited leptons the observed and expected limits are same in both ATLAS and CMS.

result in a higher acceptance for  $\geq 3$  charged lepton final states. A similar search for excited leptons with  $l^* \rightarrow l + \gamma$  decays ( $l = e, \mu$ ) produced in contact interactions is performed by the CMS Collaboration at  $\sqrt{s} = 7$  TeV [80], resulting in a mass exclusion of 1.9 TeV for  $\Lambda = m_{l^*}$ . Figure 4 summarizes the most stringent 95% C.L. lower mass limits for the excited quarks and leptons obtained from the LHC experiments.

### References

1. E.J. Eichten, K.D. Lane, and M.E. Peskin, Phys. Rev. Lett. **50**, 811 (1983).
2. E.J. Eichten *et al.*, Rev. Mod. Phys. **56**, 579 (1984); Erratum *ibid.* **58**, 1065 (1986).
3. J. Gao *et al.*, Phys. Rev. Lett. **106**, 142001 (2011).
4. G. Arnison *et al.* [UA1 Collab.], Phys. Lett. **B177**, 244 (1986).
5. J.A. Appel *et al.* [UA2 Collab.], Phys. Lett. **B160**, 349 (1985).
6. F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **62**, 613 (1989);  
F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **69**, 2896 (1992);  
F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **77**, 5336 (1996);  
F. Abe *et al.* [CDF Collab.], Erratum Phys. Rev. Lett. **78**, 4307 (1997).
7. B. Abbott *et al.* [DØ Collab.], Phys. Rev. Lett. **80**, 666 (1998);  
B. Abbott *et al.* [DØ Collab.], Phys. Rev. Lett. **82**, 2457 (1999);  
B. Abbott *et al.* [DØ Collab.], Phys. Rev. **D62**, 031101 (2000);  
B. Abbott *et al.* [DØ Collab.], Phys. Rev. **D64**, 032003 (2001);  
B. Abbott *et al.* [DØ Collab.], Phys. Rev. Lett. **103**, 191803 (2009).
8. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. **B694**, 327 (2011);  
G. Aad *et al.* [ATLAS Collab.], New J. Phys. **13**, 053044 (2011);  
G. Aad *et al.* [ATLAS Collab.], JHEP **1301**, 029 (2013).
9. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. Lett. **114**, 221802 (2015).
10. V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. Lett. **105**, 262001 (2010);  
V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. Lett. **106**, 201804 (2011);  
S. Chatrchyan *et al.* [CMS Collab.], JHEP **1205**, 055 (2012).
11. S. Chatrchyan *et al.* [CMS Collab.], Phys. Rev. **D87**, 052017 (2013).
12. V. Khachatryan *et al.* [CMS Collab.], Phys. Lett. **B746**, 79 (2015).
13. T. Sjöstrand, S. Mrenna, and P. Skands, Comp. Phys. Comm. **178**, 852 (2008).
14. S. Agostinelli *et al.* [GEANT4 Collab.], GEANT4: a simulation toolkit, Nucl. Instrum. Methods **A506**, 250 (2003).
15. Z. Nagy, Phys. Rev. Lett. **88**, 122003 (2002);  
Z. Nagy, Phys. Rev. **D68**, 094002, (2003).
16. S. Dittmaier, A. Huss, and C. Speckner, JHEP **1211**, 095 (2012).
17. T. Junk, Nucl. Instrum. Methods **A434**, 435 (1999).
18. A. L. Read, J. Phys. **G28**, 2693 (2002).
19. S. Schael *et al.* [ALEPH Collab.], Eur. Phys. J. **C49**, 411 (2007).
20. J. Abdallah *et al.* [DELPHI Collab.], Eur. Phys. J. **C45**, 589 (2006).
21. M. Acciarri *et al.* [L3 Collab.], Phys. Lett. **B489**, 81 (2000).
22. K. Ackerstaff *et al.* [OPAL Collab.], Phys. Lett. **B391**, 221 (1997).
23. G. Abbiendi *et al.* [OPAL Collab.], Eur. Phys. J. **C33**, 173 (2004).
24. F.D. Aaron *et al.* [H1 Collab.], Phys. Lett. **B705**, 52 (2011).
25. S. Chekanov *et al.* [ZEUS Collab.], Phys. Lett. **B591**, 23 (2004).
26. F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **68**, 1463 (1992).
27. F. Abe *et al.* [CDF Collab.], Phys. Rev. Lett. **79**, 2198 (1997).
28. T. Affolder *et al.* [CDF Collab.], Phys. Rev. Lett. **87**, 231803 (2001).
29. A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **96**, 211801 (2006).
30. B. Abbott *et al.* [DØ Collab.], Phys. Rev. Lett. **82**, 4769 (1999).



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31. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. **D84**, 011101 (2011).
  32. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. **B712**, 40 (2012).
  33. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. **D87**, 015010 (2013).
  34. G. Aad *et al.* [ATLAS Collab.], Eur. Phys. J. **C74**, 3134 (2014).
  35. S. Chatrchyan *et al.* [CMS Collab.], Phys. Rev. **D87**, 032001 (2013).
  36. W. Buchmuller and D. Wyler, Nucl. Phys. **B268**, 621 (1986).
  37. B. Grzadkowski *et al.*, JHEP **1010**, 085 (2010).
  38. F.M. Renard, Phys. Lett. **B116**, 264 (1982).
  39. F. del Aguila, A. Mendez, and R. Pascual, Phys. Lett. **B140**, 431 (1984).
  40. M. Suzuki, Phys. Lett. **B143**, 237 (1984).
  41. U. Baur, M. Spira, and P.M. Zerwas, Phys. Rev. **D42**, 815 (1990).
  42. K. Hagiwara, D. Zeppenfeld, and S. Komamiya, Z. Phys. **C29**, 115 (1985).
  43. N. Cabibbo, L. Maiani, and Y. Srivastava, Phys. Lett. **B139**, 459 (1984).
  44. D. Decamp *et al.* [ALEPH Collab.], Phys. Reports **216**, 253 (1992).
  45. P. Barate *et al.* [ALEPH Collab.], Eur. Phys. J. **C4**, 571 (1998).
  46. P. Abreu *et al.* [DELPHI Collab.], Nucl. Phys. **B367**, 511 (1991).
  47. J. Abdallah *et al.* [DELPHI Collab.], Eur. Phys. J. **C37**, 405 (2004).
  48. O. Adriani *et al.* [L3 Collab.], Phys. Reports **236**, 1 (1993).
  49. P. Achard *et al.* [L3 Collab.], Phys. Lett. **B531**, 28 (2002).
  50. P. Achard *et al.* [L3 Collab.], Phys. Lett. **B568**, 23 (2003).
  51. G. Abbiendi *et al.* [OPAL Collab.], Phys. Lett. **B544**, 57 (2002).
  52. G. Abbiendi *et al.* [OPAL Collab.], Phys. Lett. **B602**, 167 (2004).
  53. C. Adloff *et al.* [H1 Collab.], Phys. Lett. **B525**, 9 (2002).
  54. F.D. Aaron *et al.* [H1 Collab.], Phys. Lett. **B663**, 382 (2008).
  55. F.D. Aaron *et al.* [H1 Collab.], Phys. Lett. **B666**, 131 (2008).
  56. S. Chekanov *et al.* [ZEUS Collab.], Phys. Lett. **B549**, 32 (2002).
  57. D. Acosta *et al.* [CDF Collab.], Phys. Rev. Lett. **94**, 101802 (2005).
  58. A. Abulencia *et al.* [CDF Collab.], Phys. Rev. Lett. **97**, 191802 (2006).
  59. T. Aaltonen *et al.* [CDF Collab.], Phys. Rev. **D79**, 112002 (2009).
  60. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D73**, 111102 (2006).
  61. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. **D77**, 091102 (2008).
  62. V.M. Abazov *et al.* [DØ Collab.], Phys. Rev. Lett. **103**, 191803 (2009).
  63. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. **B708**, 37 (2012).
  64. G. Aad *et al.* [ATLAS Collab.], JHEP **1301**, 29 (2013).
  65. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. **D91**, 052007 (2015).
  66. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. Lett. **108**, 211802 (2012).
  67. G. Aad *et al.* [ATLAS Collab.], Phys. Lett. **B728**, 562 (2014).
  68. G. Aad *et al.* [ATLAS Collab.], Phys. Rev. **D85**, 072003 (2012).
  69. G. Aad *et al.* [ATLAS Collab.], New J. Phys. **15**, 093011 (2013).
  70. G. Aad *et al.* [ATLAS Collab.], JHEP **1508**, 138 (2015).
  71. S. Chatrchyan *et al.* [CMS Collab.], Phys. Lett. **B704**, 123 (2011).
  72. S. Chatrchyan *et al.* [CMS Collab.], JHEP **1301**, 13 (2013).
  73. S. Chatrchyan *et al.* [CMS Collab.], Phys. Rev. **D87**, 114015 (2013).
  74. V. Khachatryan *et al.* [CMS Collab.], Phys. Rev. **D91**, 052009 (2015).
  75. V. Khachatryan *et al.* [CMS Collab.], Phys. Lett. **B738**, 274 (2014).
  76. S. Chatrchyan *et al.* [CMS Collab.], Phys. Lett. **B722**, 28 (2013).
  77. S. Chatrchyan *et al.* [CMS Collab.], Phys. Lett. **B723**, 280 (2013).
  78. V. Khachatryan *et al.* [CMS Collab.], JHEP **1408**, 173 (2014).
  79. S. Chatrchyan *et al.* [CMS Collab.], Phys. Lett. **B704**, 143 (2011).
  80. S. Chatrchyan *et al.* [CMS Collab.], Phys. Lett. **B720**, 309 (2013).
  81. M. Cacciari, G.P. Salam, and G. Soyez, JHEP **0804**, 063 (2008).
  82. T. Sjöstrand *et al.*, Comp. Phys. Comm. **135**, 238 (2001).
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# Searches Particle Listings

## Quark and Lepton Compositeness

### SCALE LIMITS for Contact Interactions: $\Lambda(eee)$

Limits are for  $\Lambda_{LL}^{\pm}$  only. For other cases, see each reference.

$\Lambda_{LL}^{+}(\text{TeV})$	$\Lambda_{LL}^{-}(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
> <b>8.3</b>	> <b>10.3</b>	95	<sup>1</sup> BOURILKOV 01	RVUE	$E_{\text{cm}} = 192\text{--}208\text{ GeV}$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •			
>4.5	>7.0	95	<sup>2</sup> SCHAEEL 07A	ALEP	$E_{\text{cm}} = 189\text{--}209\text{ GeV}$
>5.3	>6.8	95	ABDALLAH 06C	DLPH	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
>4.7	>6.1	95	<sup>3</sup> ABBIENDI 04G	OPAL	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
>4.3	>4.9	95	ACCIARRI 00P	L3	$E_{\text{cm}} = 130\text{--}189\text{ GeV}$

<sup>1</sup> A combined analysis of the data from ALEPH, DELPHI, L3, and OPAL.  
<sup>2</sup> SCHAEEL 07A limits are from  $R_C$ ,  $Q_{FB}^{\text{depl}}$ , and hadronic cross section measurements.  
<sup>3</sup> ABBIENDI 04G limits are from  $e^{+}e^{-} \rightarrow e^{+}e^{-}$  cross section at  $\sqrt{s} = 130\text{--}207\text{ GeV}$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(ee\mu)$

Limits are for  $\Lambda_{LL}^{\pm}$  only. For other cases, see each reference.

$\Lambda_{LL}^{+}(\text{TeV})$	$\Lambda_{LL}^{-}(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
>6.6	> <b>9.5</b>	95	<sup>1</sup> SCHAEEL 07A	ALEP	$E_{\text{cm}} = 189\text{--}209\text{ GeV}$
> <b>8.5</b>	>3.8	95	ACCIARRI 00P	L3	$E_{\text{cm}} = 130\text{--}189\text{ GeV}$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •			
>7.3	>7.6	95	ABDALLAH 06C	DLPH	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
>8.1	>7.3	95	<sup>2</sup> ABBIENDI 04G	OPAL	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$

<sup>1</sup> SCHAEEL 07A limits are from  $R_C$ ,  $Q_{FB}^{\text{depl}}$ , and hadronic cross section measurements.  
<sup>2</sup> ABBIENDI 04G limits are from  $e^{+}e^{-} \rightarrow \mu\mu$  cross section at  $\sqrt{s} = 130\text{--}207\text{ GeV}$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(e\tau\tau)$

Limits are for  $\Lambda_{LL}^{\pm}$  only. For other cases, see each reference.

$\Lambda_{LL}^{+}(\text{TeV})$	$\Lambda_{LL}^{-}(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
> <b>7.9</b>	>5.8	95	<sup>1</sup> SCHAEEL 07A	ALEP	$E_{\text{cm}} = 189\text{--}209\text{ GeV}$
> <b>7.9</b>	>4.6	95	ABDALLAH 06C	DLPH	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
>4.9	> <b>7.2</b>	95	<sup>2</sup> ABBIENDI 04G	OPAL	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •			
>5.4	>4.7	95	ACCIARRI 00P	L3	$E_{\text{cm}} = 130\text{--}189\text{ GeV}$

<sup>1</sup> SCHAEEL 07A limits are from  $R_C$ ,  $Q_{FB}^{\text{depl}}$ , and hadronic cross section measurements.  
<sup>2</sup> ABBIENDI 04G limits are from  $e^{+}e^{-} \rightarrow \tau\tau$  cross section at  $\sqrt{s} = 130\text{--}207\text{ GeV}$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(\ell\ell\ell\ell)$

Lepton universality assumed. Limits are for  $\Lambda_{LL}^{\pm}$  only. For other cases, see each reference.

$\Lambda_{LL}^{+}(\text{TeV})$	$\Lambda_{LL}^{-}(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
>7.9	> <b>10.3</b>	95	<sup>1</sup> SCHAEEL 07A	ALEP	$E_{\text{cm}} = 189\text{--}209\text{ GeV}$
> <b>9.1</b>	>8.2	95	ABDALLAH 06C	DLPH	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •			
>7.7	>9.5	95	<sup>2</sup> ABBIENDI 04G	OPAL	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
>9.0	>5.2	95	<sup>3</sup> BABICH 03	RVUE	$E_{\text{cm}} = 130\text{--}207\text{ GeV}$
		95	ACCIARRI 00P	L3	$E_{\text{cm}} = 130\text{--}189\text{ GeV}$

<sup>1</sup> SCHAEEL 07A limits are from  $R_C$ ,  $Q_{FB}^{\text{depl}}$ , and hadronic cross section measurements.  
<sup>2</sup> ABBIENDI 04G limits are from  $e^{+}e^{-} \rightarrow \ell^{+}\ell^{-}$  cross section at  $\sqrt{s} = 130\text{--}207\text{ GeV}$ .  
<sup>3</sup> BABICH 03 obtain a bound  $-0.175\text{ TeV}^{-2} < 1/\Lambda_{LL}^2 < 0.095\text{ TeV}^{-2}$  (95%CL) in a model independent analysis allowing all of  $\Lambda_{LL}$ ,  $\Lambda_{LR}$ ,  $\Lambda_{RL}$ ,  $\Lambda_{RR}$  to coexist.

### SCALE LIMITS for Contact Interactions: $\Lambda(eeqq)$

Limits are for  $\Lambda_{LL}^{\pm}$  only. For other cases, see each reference.

$\Lambda_{LL}^{+}(\text{TeV})$	$\Lambda_{LL}^{-}(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
> <b>16.4</b>	> <b>20.7</b>	95	<sup>1</sup> AAD 14BE	ATLS	( $eeqq$ )
> 8.4	> <b>10.2</b>	95	<sup>2</sup> ABDALLAH 09	DLPH	( $eebb$ )
> <b>9.4</b>	> <b>5.6</b>	95	<sup>3</sup> SCHAEEL 07A	ALEP	( $eecc$ )
> <b>9.4</b>	>4.9	95	<sup>2</sup> SCHAEEL 07A	ALEP	( $eebb$ )
> <b>23.3</b>	> <b>12.5</b>	95	<sup>4</sup> CHEUNG 01B	RVUE	( $eeuu$ )
> <b>11.1</b>	> <b>26.4</b>	95	<sup>4</sup> CHEUNG 01B	RVUE	( $eedd$ )
• • •	We do not use the following data for averages, fits, limits, etc.	• • •			
>13.5	>18.3	95	<sup>5</sup> KHACHATRY...15AE	CMS	( $eeqq$ )
> 9.5	>12.1	95	<sup>6</sup> AAD 13E	ATLS	( $eeqq$ )
>10.1	>9.4	95	<sup>7</sup> AAD 12AB	ATLS	( $eeqq$ )
> 4.2	>4.0	95	<sup>8</sup> AARON 11C	H1	( $eeqq$ )
> 3.8	>3.8	95	<sup>9</sup> ABDALLAH 11	DLPH	( $eetc$ )
>12.9	>7.2	95	<sup>10</sup> SCHAEEL 07A	ALEP	( $eeqq$ )
> 3.7	>5.9	95	<sup>11</sup> ABULENCIA 06L	CDF	( $eeqq$ )

- <sup>1</sup> AAD 14BE limits are from  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$ . The quoted limit uses a uniform positive prior in  $1/\Lambda^2$ .  
<sup>2</sup> ABDALLAH 09 and SCHAEEL 07A limits are from  $R_b$ ,  $A_{FB}^b$ .  
<sup>3</sup> SCHAEEL 07A limits are from  $R_C$ ,  $Q_{FB}^{\text{depl}}$ , and hadronic cross section measurements.  
<sup>4</sup> CHEUNG 01B is an update of BARGER 98E.  
<sup>5</sup> KHACHATRYAN 15AE limit is from  $e^{+}e^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ .  
<sup>6</sup> AAD 13E limits are from  $e^{+}e^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ .  
<sup>7</sup> AAD 12AB limits are from  $e^{+}e^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ .  
<sup>8</sup> AARON 11C limits are from  $Q^2$  spectrum measurements of  $e^{\pm}p \rightarrow e^{\pm}X$ .  
<sup>9</sup> ABDALLAH 11 limit is from  $e^{+}e^{-} \rightarrow t\bar{c}$  cross section.  $\Lambda_{LL} = \Lambda_{LR} = \Lambda_{RL} = \Lambda_{RR}$  is assumed.  
<sup>10</sup> SCHAEEL 07A limit assumes quark flavor universality of the contact interactions.  
<sup>11</sup> ABULENCIA 06L limits are from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96\text{ TeV}$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(\mu\mu qq)$

$\Lambda_{LL}^{+}(\text{TeV})$	$\Lambda_{LL}^{-}(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
> <b>12.5</b>	> <b>16.7</b>	95	<sup>1</sup> AAD 14BE	ATLS	( $\mu\mu qq$ )
• • •	We do not use the following data for averages, fits, limits, etc.	• • •			
>12.0	>15.2	95	<sup>2</sup> KHACHATRY...15AE	CMS	( $\mu\mu qq$ )
> 9.6	>12.9	95	<sup>3</sup> AAD 13E	ATLS	( $\mu\mu qq$ ) (isosinglet)
> 9.5	>13.1	95	<sup>4</sup> CHATRCHYAN13K	CMS	( $\mu\mu qq$ ) (isosinglet)
> 8.0	>7.0	95	<sup>5</sup> AAD 12AB	ATLS	( $\mu\mu qq$ ) (isosinglet)

<sup>1</sup> AAD 14BE limits are from  $pp$  collisions at  $\sqrt{s} = 8\text{ TeV}$ . The quoted limit uses a uniform positive prior in  $1/\Lambda^2$ .  
<sup>2</sup> KHACHATRYAN 15AE limit is from  $\mu^{+}\mu^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ .  
<sup>3</sup> AAD 13E limits are from  $\mu^{+}\mu^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ .  
<sup>4</sup> CHATRCHYAN 13K limits are from  $\mu^{+}\mu^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ .  
<sup>5</sup> AAD 12AB limits are from  $\mu^{+}\mu^{-}$  mass distribution in  $pp$  collisions at  $E_{\text{cm}} = 7\text{ TeV}$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(\ell\nu\ell\nu)$

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
> <b>3.10</b>	90	<sup>1</sup> JODIDIO 86	SPEC	$\Lambda_{LR}^{\pm}(\nu_{\mu}\nu_{\mu}e\mu e)$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
>3.8		<sup>2</sup> DIAZCRUZ 94	RVUE	$\Lambda_{LL}^{+}(\tau\nu_{\tau}e\nu_e)$
>8.1		<sup>2</sup> DIAZCRUZ 94	RVUE	$\Lambda_{LL}^{-}(\tau\nu_{\tau}e\nu_e)$
>4.1		<sup>3</sup> DIAZCRUZ 94	RVUE	$\Lambda_{LL}^{+}(\tau\nu_{\tau}\mu\nu_{\mu})$
>6.5		<sup>3</sup> DIAZCRUZ 94	RVUE	$\Lambda_{LL}^{-}(\tau\nu_{\tau}\mu\nu_{\mu})$

<sup>1</sup> JODIDIO 86 limit is from  $\mu^{+} \rightarrow \bar{\nu}_{\mu}e^{+}\nu_e$ . Chirality invariant interactions  $L = (g^2/\Lambda^2)$  [ $\eta_{LL}(\bar{\nu}_{\mu}L\gamma^{\alpha}\mu_L)(\bar{e}L\gamma_{\alpha}\nu_eL) + \eta_{LR}(\bar{\nu}_{\mu}L\gamma^{\alpha}\nu_eL)(\bar{e}R\gamma_{\alpha}\mu_R)$ ] with  $g^2/4\pi = 1$  and  $(\eta_{LL}, \eta_{LR}) = (0, \pm 1)$  are taken. No limits are given for  $\Lambda_{LL}^{\pm}$  with  $(\eta_{LL}, \eta_{LR}) = (\pm 1, 0)$ . For more general constraints with right-handed neutrinos and chirality nonconserving contact interactions, see their text.  
<sup>2</sup> DIAZCRUZ 94 limits are from  $\Gamma(\tau \rightarrow e\nu\nu)$  and assume flavor-dependent contact interactions with  $\Lambda(\tau\nu_{\tau}e\nu_e) \ll \Lambda(\mu\nu_{\mu}e\nu_e)$ .  
<sup>3</sup> DIAZCRUZ 94 limits are from  $\Gamma(\tau \rightarrow \mu\nu\nu)$  and assume flavor-dependent contact interactions with  $\Lambda(\tau\nu_{\tau}\mu\nu_{\mu}) \ll \Lambda(\mu\nu_{\mu}e\nu_e)$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(e\nu qq)$

VALUE (TeV)	CL%	DOCUMENT ID	TECN
> <b>2.81</b>	95	<sup>1</sup> AFFOLDER 00I	CDF

<sup>1</sup> AFFOLDER 00I bound is for a scalar interaction  $\bar{q}_R q_L \bar{\nu}_e L$ .

### SCALE LIMITS for Contact Interactions: $\Lambda(qqqq)$

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
>8.1	95	<sup>1</sup> AAD 15L	ATLS	$pp$ dijet angl. $\Lambda_{LL}^{+}$
> <b>9.0</b>	95	<sup>2</sup> KHACHATRY...15J	CMS	$pp$ dijet angl. $\Lambda_{LL}^{+}$
• • •	We do not use the following data for averages, fits, limits, etc.	• • •		
>5	95	<sup>3</sup> AAD 15BY	ATLS	$pp \rightarrow t\bar{t}t\bar{t}$
>7.6	95	<sup>4</sup> FABBRICHESI 14	RVUE	$q\bar{q}t\bar{t}$
>9.9	95	<sup>5</sup> AAD 13D	ATLS	$pp \rightarrow$ dijet angl.
>7.5	95	<sup>6</sup> CHATRCHYAN13AN	CMS	$pp \rightarrow$ dijet.; $\Lambda_{LL}^{+}$
		<sup>7</sup> CHATRCHYAN12Z	CMS	$pp \rightarrow$ dijet angl.; $\Lambda_{LL}^{+}$

<sup>1</sup> AAD 15L limit is from dijet angular distribution in  $pp$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ .  $u$ ,  $d$ , and  $s$  quarks are assumed to be composite. They also obtain  $\Lambda_{LL}^{-} > 12.0\text{ TeV}$ .  
<sup>2</sup> KHACHATRYAN 15J limit is from dijet angular distribution in  $pp$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ .  $u$ ,  $d$ ,  $s$ ,  $c$ , and  $b$  quarks are assumed to be composite. They also obtain  $\Lambda_{LL}^{-} > 11.7\text{ TeV}$ .  
<sup>3</sup> AAD 15BY obtain limit on the  $t_R$  compositeness  $2\pi\Lambda_{RR}^2 < 15.1\text{ TeV}^{-2}$  at 95% CL from the  $t\bar{t}t\bar{t}$  production in the  $pp$  collisions at  $E_{\text{cm}} = 8\text{ TeV}$ .  
<sup>4</sup> FABBRICHESI 14 obtain bounds on chromoelectric and chromomagnetic form factors of the top-quark using  $pp \rightarrow t\bar{t}$  and  $p\bar{p} \rightarrow t\bar{t}$  cross sections. The quoted limit on the  $q\bar{q}t\bar{t}$  contact interaction is derived from their bound on the chromoelectric form factor.

See key on page 601

## Searches Particle Listings

### Quark and Lepton Compositeness

<sup>5</sup> AAD 13b limit is from dijet angular distribution in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. The constant prior in  $1/\Lambda^4$  is applied.

<sup>6</sup> CHATRCHYAN 13AN limit is from inclusive jet  $p_T$  spectrum in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. They also obtain  $\Lambda_{LL}^- > 14.3$  TeV.

<sup>7</sup> CHATRCHYAN 12Z limit is from dijet angular distribution in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV. They also obtain  $\Lambda_{LL}^- > 10.5$  TeV.

#### SCALE LIMITS for Contact Interactions: $\Lambda(\nu\nu qq)$

Limits are for  $\Lambda_{LL}^{\pm}$  only. For other cases, see each reference.

$\Lambda_{LL}^+(\text{TeV})$	$\Lambda_{LL}^-(\text{TeV})$	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;5.0</b>	<b>&gt;5.4</b>	95	<sup>1</sup> MCFARLAND 98	CCFR	$\nu N$ scattering

<sup>1</sup> MCFARLAND 98 assumed a flavor universal interaction. Neutrinos were mostly of muon type.

#### MASS LIMITS for Excited $e$ ( $e^*$ )

Most  $e^+e^-$  experiments assume one-photon or  $Z$  exchange. The limits from some  $e^+e^-$  experiments which depend on  $\lambda$  have assumed transition couplings which are chirality violating ( $\eta_L = \eta_R$ ). However they can be interpreted as limits for chirality-conserving interactions after multiplying the coupling value  $\lambda$  by  $\sqrt{2}$ ; see Note.

Excited leptons have the same quantum numbers as other ortho-leptons. See also the searches for ortho-leptons in the "Searches for Heavy Leptons" section.

#### Limits for Excited $e$ ( $e^*$ ) from Pair Production

These limits are obtained from  $e^+e^- \rightarrow e^{*+}e^{*-}$  and thus rely only on the (electroweak) charge of  $e^*$ . Form factor effects are ignored unless noted. For the case of limits from  $Z$  decay, the  $e^*$  coupling is assumed to be of sequential type. Possible  $t$ -channel contribution from transition magnetic coupling is neglected. All limits assume a dominant  $e^* \rightarrow e\gamma$  decay except the limits from  $\Gamma(Z)$ .

For limits prior to 1987, see our 1992 edition (Physical Review **D45** S1 (1992)).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;103.2</b>	95	<sup>1</sup> ABBIENDI 02G	OPAL	$e^+e^- \rightarrow e^*e^*$ Homodoublet type
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>102.8	95	<sup>2</sup> ACHARD 03B	L3	$e^+e^- \rightarrow e^*e^*$ Homodoublet type

<sup>1</sup> From  $e^+e^-$  collisions at  $\sqrt{s} = 183$ –209 GeV.  $f = f'$  is assumed.

<sup>2</sup> From  $e^+e^-$  collisions at  $\sqrt{s} = 189$ –209 GeV.  $f = f'$  is assumed. ACHARD 03B also obtain limit for  $f = -f'$ :  $m_{e^*} > 96.6$  GeV.

#### Limits for Excited $e$ ( $e^*$ ) from Single Production

These limits are from  $e^+e^- \rightarrow e^*e$ ,  $W \rightarrow e^*\nu$ , or  $ep \rightarrow e^*X$  and depend on transition magnetic coupling between  $e$  and  $e^*$ . All limits assume  $e^* \rightarrow e\gamma$  decay except as noted. Limits from LEP, UA2, and H1 are for chiral coupling, whereas all other limits are for nonchiral coupling,  $\eta_L = \eta_R = 1$ . In most papers, the limit is expressed in the form of an excluded region in the  $\lambda$ – $m_{e^*}$  plane. See the original papers.

For limits prior to 1987, see our 1992 edition (Physical Review **D45** S1 (1992)).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;3000</b>	95	<sup>1</sup> AAD 15AP	ATLS	$pp \rightarrow e^{(*)}e^*X$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>2200	95	<sup>2</sup> AAD 13BB	ATLS	$pp \rightarrow ee^*X$
>1900	95	<sup>3</sup> CHATRCHYAN 13AE	CMS	$pp \rightarrow ee^*X$
>1870	95	<sup>4</sup> AAD 12AZ	ATLS	$pp \rightarrow e^{(*)}e^*X$
>1070	95	<sup>5</sup> CHATRCHYAN 11X	CMS	$pp \rightarrow ee^*X$

<sup>1</sup> AAD 15AP search for  $e^*$  production in evens with three or more charged leptons in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The quoted limit assumes  $\Lambda = m_{e^*}$ ,  $f = f' = 1$ . The contact interaction is included in the  $e^*$  production and decay amplitudes.

<sup>2</sup> AAD 13BB search for single  $e^*$  production in  $pp$  collisions with  $e^* \rightarrow e\gamma$  decay.  $f = f' = 1$ , and  $e^*$  production via contact interaction with  $\Lambda = m_{e^*}$  are assumed.

<sup>3</sup> CHATRCHYAN 13AE search for single  $e^*$  production in  $pp$  collisions with  $e^* \rightarrow e\gamma$  decay.  $f = f' = 1$ , and  $e^*$  production via contact interaction with  $\Lambda = m_{e^*}$  are assumed.

<sup>4</sup> AAD 12AZ search for  $e^*$  production via four-fermion contact interaction in  $pp$  collisions with  $e^* \rightarrow e\gamma$  decay. The quoted limit assumes  $\Lambda = m_{e^*}$ . See their Fig. 8 for the exclusion plot in the mass-coupling plane.

<sup>5</sup> CHATRCHYAN 11X search for single  $e^*$  production in  $pp$  collisions with the decay  $e^* \rightarrow e\gamma$ .  $f = f' = \Lambda/m_{e^*}$  is assumed. See their Fig. 2 for the exclusion plot in the mass-coupling plane.

#### Limits for Excited $e$ ( $e^*$ ) from $e^+e^- \rightarrow \gamma\gamma$

These limits are derived from indirect effects due to  $e^*$  exchange in the  $t$  channel and depend on transition magnetic coupling between  $e$  and  $e^*$ . All limits are for  $\lambda_\gamma = 1$ . All limits except ABE 89J and ACHARD 02D are for nonchiral coupling with  $\eta_L = \eta_R = 1$ . We choose the chiral coupling limit as the best limit and list it in the Summary Table.

For limits prior to 1987, see our 1992 edition (Physical Review **D45** S1 (1992)).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;356</b>	95	<sup>1</sup> ABDALLAH 04N	DLPH	$\sqrt{s} = 161$ –208 GeV
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>310	95	ACHARD 02D	L3	$\sqrt{s} = 192$ –209 GeV

<sup>1</sup> ABDALLAH 04N also obtain a limit on the excited electron mass with  $e^*e^*$  chiral coupling,  $m_{e^*} > 295$  GeV at 95% CL.

#### Indirect Limits for Excited $e$ ( $e^*$ )

These limits make use of loop effects involving  $e^*$  and are therefore subject to theoretical uncertainty.

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
• • •	We do not use the following data for averages, fits, limits, etc. • • •		
<sup>1</sup> DORENBOS... 89	CHRM	$\overline{\nu}_\mu e \rightarrow \overline{\nu}_\mu e, \nu_\mu e \rightarrow \nu_\mu e$	
<sup>2</sup> GRIFOLS 86	THEO	$\nu_\mu e \rightarrow \nu_\mu e$	
<sup>3</sup> RENARD 82	THEO	$g-2$ of electron	

<sup>1</sup> DORENBOSCH 89 obtain the limit  $\lambda_\gamma^2 \Lambda_{\text{cut}}^2 / m_{e^*}^2 < 2.6$  (95% CL), where  $\Lambda_{\text{cut}}$  is the cutoff scale, based on the one-loop calculation by GRIFOLS 86. If one assumes that  $\Lambda_{\text{cut}} = 1$  TeV and  $\lambda_\gamma = 1$ , one obtains  $m_{e^*} > 620$  GeV. However, one generally expects  $\lambda_\gamma \approx m_{e^*} / \Lambda_{\text{cut}}$  in composite models.

<sup>2</sup> GRIFOLS 86 uses  $\nu_\mu e \rightarrow \nu_\mu e$  and  $\overline{\nu}_\mu e \rightarrow \overline{\nu}_\mu e$  data from CHARM Collaboration to derive mass limits which depend on the scale of compositeness.

<sup>3</sup> RENARD 82 derived from  $g-2$  data limits on mass and couplings of  $e^*$  and  $\mu^*$ . See figures 2 and 3 of the paper.

#### MASS LIMITS for Excited $\mu$ ( $\mu^*$ )

#### Limits for Excited $\mu$ ( $\mu^*$ ) from Pair Production

These limits are obtained from  $e^+e^- \rightarrow \mu^{*+}\mu^{*-}$  and thus rely only on the (electroweak) charge of  $\mu^*$ . Form factor effects are ignored unless noted. For the case of limits from  $Z$  decay, the  $\mu^*$  coupling is assumed to be of sequential type. All limits assume a dominant  $\mu^* \rightarrow \mu\gamma$  decay except the limits from  $\Gamma(Z)$ .

For limits prior to 1987, see our 1992 edition (Physical Review **D45** S1 (1992)).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;103.2</b>	95	<sup>1</sup> ABBIENDI 02G	OPAL	$e^+e^- \rightarrow \mu^*\mu^*$ Homodoublet type
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>102.8	95	<sup>2</sup> ACHARD 03B	L3	$e^+e^- \rightarrow \mu^*\mu^*$ Homodoublet type

<sup>1</sup> From  $e^+e^-$  collisions at  $\sqrt{s} = 183$ –209 GeV.  $f = f'$  is assumed.

<sup>2</sup> From  $e^+e^-$  collisions at  $\sqrt{s} = 189$ –209 GeV.  $f = f'$  is assumed. ACHARD 03B also obtain limit for  $f = -f'$ :  $m_{\mu^*} > 96.6$  GeV.

#### Limits for Excited $\mu$ ( $\mu^*$ ) from Single Production

These limits are from  $e^+e^- \rightarrow \mu^*\mu$  and depend on transition magnetic coupling between  $\mu$  and  $\mu^*$ . All limits assume  $\mu^* \rightarrow \mu\gamma$  decay. Limits from LEP are for chiral coupling, whereas all other limits are for nonchiral coupling,  $\eta_L = \eta_R = 1$ . In most papers, the limit is expressed in the form of an excluded region in the  $\lambda$ – $m_{\mu^*}$  plane. See the original papers.

For limits prior to 1987, see our 1992 edition (Physical Review **D45** S1 (1992)).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;3000</b>	95	<sup>1</sup> AAD 15AP	ATLS	$pp \rightarrow \mu^{(*)}\mu^*X$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>2200	95	<sup>2</sup> AAD 13BB	ATLS	$pp \rightarrow \mu\mu^*X$
>1900	95	<sup>3</sup> CHATRCHYAN 13AE	CMS	$pp \rightarrow \mu\mu^*X$
>1750	95	<sup>4</sup> AAD 12AZ	ATLS	$pp \rightarrow \mu^{(*)}\mu^*X$
>1090	95	<sup>5</sup> CHATRCHYAN 11X	CMS	$pp \rightarrow \mu\mu^*X$

<sup>1</sup> AAD 15AP search for  $\mu^*$  production in evens with three or more charged leptons in  $pp$  collisions at  $\sqrt{s} = 8$  TeV. The quoted limit assumes  $\Lambda = m_{\mu^*}$ ,  $f = f' = 1$ . The contact interaction is included in the  $\mu^*$  production and decay amplitudes.

<sup>2</sup> AAD 13BB search for single  $\mu^*$  production in  $pp$  collisions with  $\mu^* \rightarrow \mu\gamma$  decay.  $f = f' = 1$ , and  $\mu^*$  production via contact interaction with  $\Lambda = m_{\mu^*}$  are assumed.

<sup>3</sup> CHATRCHYAN 13AE search for single  $\mu^*$  production in  $pp$  collisions with  $\mu^* \rightarrow \mu\gamma$  decay.  $f = f' = 1$ , and  $\mu^*$  production via contact interaction with  $\Lambda = m_{\mu^*}$  are assumed.

<sup>4</sup> AAD 12AZ search for  $\mu^*$  production via four-fermion contact interaction in  $pp$  collisions with  $\mu^* \rightarrow \mu\gamma$  decay. The quoted limit assumes  $\Lambda = m_{\mu^*}$ . See their Fig. 8 for the exclusion plot in the mass-coupling plane.

<sup>5</sup> CHATRCHYAN 11X search for single  $\mu^*$  production in  $pp$  collisions with the decay  $\mu^* \rightarrow \mu\gamma$ .  $f = f' = \Lambda/m_{\mu^*}$  is assumed. See their Fig. 2 for the exclusion plot in the mass-coupling plane.

# Searches Particle Listings

## Quark and Lepton Compositeness

### Indirect Limits for Excited $\mu$ ( $\mu^*$ )

These limits make use of loop effects involving  $\mu^*$  and are therefore subject to theoretical uncertainty.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
		<sup>1</sup> RENARD	82	THEO $g_{-2}$ of muon
<sup>1</sup> RENARD 82 derived from $g_{-2}$ data limits on mass and couplings of $e^*$ and $\mu^*$ . See figures 2 and 3 of the paper.				

### MASS LIMITS for Excited $\tau$ ( $\tau^*$ )

#### Limits for Excited $\tau$ ( $\tau^*$ ) from Pair Production

These limits are obtained from  $e^+e^- \rightarrow \tau^{*+}\tau^{*-}$  and thus rely only on the (electroweak) charge of  $\tau^*$ . Form factor effects are ignored unless noted. For the case of limits from Z decay, the  $\tau^*$  coupling is assumed to be of sequential type. All limits assume a dominant  $\tau^* \rightarrow \tau\gamma$  decay except the limits from  $\Gamma(Z)$ .

For limits prior to 1987, see our 1992 edition (Physical Review **D45** S1 (1992)).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>103.2	95	<sup>1</sup> ABBIENDI	02G OPAL	$e^+e^- \rightarrow \tau^*\tau^*$ Homodoublet type
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>102.8	95	<sup>2</sup> ACHARD	03B L3	$e^+e^- \rightarrow \tau^*\tau^*$ Homodoublet type
<sup>1</sup> From $e^+e^-$ collisions at $\sqrt{s} = 183\text{--}209$ GeV. $f = f'$ is assumed.				
<sup>2</sup> From $e^+e^-$ collisions at $\sqrt{s} = 189\text{--}209$ GeV. $f = f'$ is assumed. ACHARD 03B also obtain limit for $f = -f'$ : $m_{\tau^*} > 96.6$ GeV.				

#### Limits for Excited $\tau$ ( $\tau^*$ ) from Single Production

These limits are from  $e^+e^- \rightarrow \tau^*\tau$  and depend on transition magnetic coupling between  $\tau$  and  $\tau^*$ . All limits assume  $\tau^* \rightarrow \tau\gamma$  decay. Limits from LEP are for chiral coupling, whereas all other limits are for nonchiral coupling,  $\eta_L = \eta_R = 1$ . In most papers, the limit is expressed in the form of an excluded region in the  $\lambda\text{--}m_{\tau^*}$  plane. See the original papers.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>2500	95	<sup>1</sup> AAD	15AP ATLS	$p\bar{p} \rightarrow \tau^{(*)}\tau^*X$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
> 180	95	<sup>2</sup> ACHARD	03B L3	$e^+e^- \rightarrow \tau\tau^*$
> 185	95	<sup>3</sup> ABBIENDI	02G OPAL	$e^+e^- \rightarrow \tau\tau^*$
<sup>1</sup> AAD 15AP search for $\tau^*$ production in events with three or more charged leptons in $p\bar{p}$ collisions at $\sqrt{s} = 8$ TeV. The quoted limit assumes $\Lambda = m_{\tau^*}$ , $f = f' = 1$ . The contact interaction is included in the $\tau^*$ production and decay amplitudes.				
<sup>2</sup> ACHARD 03B result is from $e^+e^-$ collisions at $\sqrt{s} = 189\text{--}209$ GeV. $f = f' = \Lambda/m_{\tau^*}$ is assumed. See their Fig. 4 for the exclusion plot in the mass-coupling plane.				
<sup>3</sup> ABBIENDI 02G result is from $e^+e^-$ collisions at $\sqrt{s} = 183\text{--}209$ GeV. $f = f' = \Lambda/m_{\tau^*}$ is assumed for $\tau^*$ coupling. See their Fig. 4c for the exclusion limit in the mass-coupling plane.				

### MASS LIMITS for Excited Neutrino ( $\nu^*$ )

#### Limits for Excited $\nu$ ( $\nu^*$ ) from Pair Production

These limits are obtained from  $e^+e^- \rightarrow \nu^*\nu^*$  and thus rely only on the (electroweak) charge of  $\nu^*$ . Form factor effects are ignored unless noted. The  $\nu^*$  coupling is assumed to be of sequential type unless otherwise noted. All limits assume a dominant  $\nu^* \rightarrow \nu\gamma$  decay except the limits from  $\Gamma(Z)$ .

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>1600	95	<sup>1</sup> AAD	15AP ATLS	$p\bar{p} \rightarrow \nu^*\nu^*X$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
		<sup>2</sup> ABBIENDI	04N OPAL	
> 102.6	95	<sup>3</sup> ACHARD	03B L3	$e^+e^- \rightarrow \nu^*\nu^*$ Homodoublet type
<sup>1</sup> AAD 15AP search for $\nu^*$ pair production in events with three or more charged leptons in $p\bar{p}$ collisions at $\sqrt{s} = 8$ TeV. The quoted limit assumes $\Lambda = m_{\nu^*}$ , $f = f' = 1$ . The contact interaction is included in the $\nu^*$ production and decay amplitudes.				
<sup>2</sup> From $e^+e^-$ collisions at $\sqrt{s} = 192\text{--}209$ GeV, ABBIENDI 04N obtain limit on $\sigma(e^+e^- \rightarrow \nu^*\nu^*) B_2^2(\nu^* \rightarrow \nu\gamma)$ . See their Fig.2. The limit ranges from 20 to 45 fb for $m_{\nu^*} > 45$ GeV.				
<sup>3</sup> From $e^+e^-$ collisions at $\sqrt{s} = 189\text{--}209$ GeV. $f = -f'$ is assumed. ACHARD 03B also obtain limit for $f = f'$ : $m_{\nu_e^*} > 101.7$ GeV, $m_{\nu_\mu^*} > 101.8$ GeV, and $m_{\nu_\tau^*} > 92.9$ GeV. See their Fig. 4 for the exclusion plot in the mass-coupling plane.				

#### Limits for Excited $\nu$ ( $\nu^*$ ) from Single Production

These limits are from  $e^+e^- \rightarrow \nu\nu^*$ ,  $Z \rightarrow \nu\nu^*$ , or  $e\bar{p} \rightarrow \nu^*X$  and depend on transition magnetic coupling between  $\nu/e$  and  $\nu^*$ . Assumptions about  $\nu^*$  decay mode are given in footnotes.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>213	95	<sup>1</sup> AARON	08 H1	$e\bar{p} \rightarrow \nu^*X$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>190	95	<sup>2</sup> ACHARD	03B L3	$e^+e^- \rightarrow \nu\nu^*$
none 50–150	95	<sup>3</sup> ADLOFF	02 H1	$e\bar{p} \rightarrow \nu^*X$
>158	95	<sup>4</sup> CHEKANOV	02D ZEUS	$e\bar{p} \rightarrow \nu^*X$

<sup>1</sup> AARON 08 search for single  $\nu^*$  production in  $e\bar{p}$  collisions with the decays  $\nu^* \rightarrow \nu\gamma$ ,  $\nu Z$ ,  $eW$ . The quoted limit assumes  $f = -f' = \Lambda/m_{\nu^*}$ . See their Fig. 3 and Fig. 4 for the exclusion plots in the mass-coupling plane.

<sup>2</sup> ACHARD 03B result is from  $e^+e^-$  collisions at  $\sqrt{s} = 189\text{--}209$  GeV. The quoted limit is for  $\nu_e^*$ .  $f = -f' = \Lambda/m_{\nu^*}$  is assumed. See their Fig. 4 for the exclusion plot in the mass-coupling plane.

<sup>3</sup> ADLOFF 02 search for single  $\nu^*$  production in  $e\bar{p}$  collisions with the decays  $\nu^* \rightarrow \nu\gamma$ ,  $\nu Z$ ,  $eW$ . The quoted limit assumes  $f = -f' = \Lambda/m_{\nu^*}$ . See their Fig. 1 for the exclusion plots in the mass-coupling plane.

<sup>4</sup> CHEKANOV 02D search for single  $\nu^*$  production in  $e\bar{p}$  collisions with the decays  $\nu^* \rightarrow \nu\gamma$ ,  $\nu Z$ ,  $eW$ .  $f = -f' = \Lambda/m_{\nu^*}$  is assumed for the  $e^*$  coupling. CHEKANOV 02D also obtain limit for  $f = f' = \Lambda/m_{\nu^*}$ :  $m_{\nu_\mu^*} > 135$  GeV. See their Fig. 5c and Fig. 5d for the exclusion plot in the mass-coupling plane.

### MASS LIMITS for Excited $q$ ( $q^*$ )

#### Limits for Excited $q$ ( $q^*$ ) from Pair Production

These limits are mostly obtained from  $e^+e^- \rightarrow q^*\bar{q}^*$  and thus rely only on the (electroweak) charge of the  $q^*$ . Form factor effects are ignored unless noted. Assumptions about the  $q^*$  decay are given in the comments and footnotes.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>338	95	<sup>1</sup> AALTONEN	10H CDF	$q^* \rightarrow tW^-$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>655	95	<sup>2</sup> AAD	14AZ ATLS	$T \rightarrow tZ$
		<sup>3</sup> BARATE	98U ALEP	$Z \rightarrow q^*q^*$
> 45.6	95	<sup>4</sup> ADRIANI	93M L3	$u$ or $d$ type, $Z \rightarrow q^*q^*$
> 41.7	95	<sup>5</sup> BARDADIN...	92 RVUE	$u$ -type, $\Gamma(Z)$
> 44.7	95	<sup>5</sup> BARDADIN...	92 RVUE	$d$ -type, $\Gamma(Z)$
> 40.6	95	<sup>6</sup> DECAMP	92 ALEP	$u$ -type, $\Gamma(Z)$
> 44.2	95	<sup>6</sup> DECAMP	92 ALEP	$d$ -type, $\Gamma(Z)$
> 45	95	<sup>7</sup> DECAMP	92 ALEP	$u$ or $d$ type, $Z \rightarrow q^*q^*$
> 45	95	<sup>6</sup> ABREU	91F DLPH	$u$ -type, $\Gamma(Z)$
> 45	95	<sup>6</sup> ABREU	91F DLPH	$d$ -type, $\Gamma(Z)$

<sup>1</sup> AALTONEN 10H obtain limits on the  $q^*q^*$  production cross section in  $p\bar{p}$  collisions. See their Fig. 3.

<sup>2</sup> AAD 14AZ quoted limit is for heavy SU(2) singlet quark  $T$ .

<sup>3</sup> BARATE 98U obtain limits on the form factor. See their Fig. 16 for limits in mass-form factor plane.

<sup>4</sup> ADRIANI 93M limit is valid for  $B(q^* \rightarrow qg) > 0.25$  (0.17) for up (down) type.

<sup>5</sup> BARDADIN-OTWINOWSKA 92 limit based on  $\Delta\Gamma(Z) < 36$  MeV.

<sup>6</sup> These limits are independent of decay modes.

<sup>7</sup> Limit is for  $B(q^* \rightarrow qg) + B(q^* \rightarrow q\gamma) = 1$ .

#### Limits for Excited $q$ ( $q^*$ ) from Single Production

These limits are from  $e^+e^- \rightarrow q^*\bar{q}$ ,  $p\bar{p} \rightarrow q^*X$ , or  $p\bar{p} \rightarrow q^*X$  and depend on transition magnetic couplings between  $q$  and  $q^*$ . Assumptions about  $q^*$  decay mode are given in the footnotes and comments.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>4060	95	<sup>1</sup> AAD	15V ATLS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qg$
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
>1390	95	<sup>2</sup> KHACHATRY...	16I CMS	$p\bar{p} \rightarrow b^*X$ , $b^* \rightarrow tW$
>3500	95	<sup>3</sup> KHACHATRY...	15V CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qg$
>3500	95	<sup>4</sup> AAD	14A ATLS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow q\gamma$
>3200	95	<sup>5</sup> KHACHATRY...	14A CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qW$
>2900	95	<sup>6</sup> KHACHATRY...	14A CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qZ$
none 700–3500	95	<sup>7</sup> KHACHATRY...	14J CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow q\gamma$
> 870	95	<sup>8</sup> AAD	13AF ATLS	$p\bar{p} \rightarrow b^*X$ , $b^* \rightarrow tW$
>1940	95	<sup>9</sup> CHATRCHYAN	13AI CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qZ, qW$
>2380	95	<sup>10</sup> CHATRCHYAN	13AJ CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qW$
>2150	95	<sup>11</sup> CHATRCHYAN	13AJ CMS	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qZ$
		<sup>12</sup> ABAZOV	11F D0	$p\bar{p} \rightarrow q^*X$ , $q^* \rightarrow qZ, qW$

<sup>1</sup> AAD 15V assume  $\Lambda = m_{q^*}$ ,  $f_5 = f = f' = 1$ . The contact interactions are not included in  $q^*$  production and decay amplitudes.

<sup>2</sup> KHACHATRYAN 16I search for  $b^*$  decaying to  $tW$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 8$  TeV.  $\kappa_L^b = g_L = 1$ ,  $\kappa_R^b = g_R = 0$  are assumed. See their Fig. 8 for limits on  $\sigma \cdot B$ .

<sup>3</sup> KHACHATRYAN 15V assume  $\Lambda = m_{q^*}$ ,  $f_5 = f = f' = 1$ . The contact interactions are not included in  $q^*$  production and decay amplitudes.

<sup>4</sup> AAD 14A assume  $\Lambda = m_{q^*}$ ,  $f_5 = f = f' = 1$ .

<sup>5</sup> KHACHATRYAN 14 use the hadronic decay of  $W$ , assuming  $\Lambda = m_{q^*}$ ,  $f_5 = f = f' = 1$ .

<sup>6</sup> KHACHATRYAN 14 use the hadronic decay of  $Z$ , assuming  $\Lambda = m_{q^*}$ ,  $f_5 = f = f' = 1$ .

<sup>7</sup> KHACHATRYAN 14J assume  $f_5 = f = f' = \Lambda / m_{q^*}$ .

<sup>8</sup> AAD 13AF search for  $b^*$  decaying to  $tW$  in  $p\bar{p}$  collisions at  $\sqrt{s} = 7$  TeV.  $\kappa_L^b = g_L = 1$ ,  $\kappa_R^b = g_R = 0$  are assumed. See their Fig.6 for limits on  $\sigma \cdot B$ .

<sup>9</sup> CHATRCHYAN 13AI assume  $q^*$  production via  $qg$  fusion and  $\Lambda = m_{q^*}$ ,  $f_5 = f = f' = 1$ .

For  $q^*$  production via  $qg$  fusion and via contact interactions, the limit becomes  $m_{q^*} > 2220$  GeV.

<sup>10</sup> CHATRCHYAN 13AJ use the hadronic decay of  $W$ .

<sup>11</sup> CHATRCHYAN 13AJ use the hadronic decay of  $Z$ .

<sup>12</sup> ABAZOV 11F search for vectorlike quarks decaying to  $W$ +jet and  $Z$ +jet in  $p\bar{p}$  collisions. See their Fig. 3 and Fig. 4 for the limits on  $\sigma \cdot B$ .

See key on page 601

Searches Particle Listings

Quark and Lepton Compositeness, Extra Dimensions

MASS LIMITS for Color Sextet Quarks ( $q_6$ )

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>84	95	<sup>1</sup> ABE	89D CDF	$p\bar{p} \rightarrow q_6\bar{q}_6$

<sup>1</sup> ABE 89D look for pair production of unit-charged particles which leave the detector before decaying. In the above limit the color sextet quark is assumed to fragment into a unit-charged or neutral hadron with equal probability and to have long enough lifetime not to decay within the detector. A limit of 121 GeV is obtained for a color decuplet.

MASS LIMITS for Color Octet Charged Leptons ( $\ell_8$ )

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>86	95	<sup>1</sup> ABE	89D CDF	Stable $\ell_8$ : $p\bar{p} \rightarrow \ell_8\bar{\ell}_8$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<sup>2</sup> ABT 93 H1  $e_8$ :  $e p \rightarrow e_8 X$

<sup>1</sup> ABE 89D look for pair production of unit-charged particles which leave the detector before decaying. In the above limit the color octet lepton is assumed to fragment into a unit-charged or neutral hadron with equal probability and to have long enough lifetime not to decay within the detector. The limit improves to 99 GeV if it always fragments into a unit-charged hadron.

<sup>2</sup> ABT 93 search for  $e_8$  production via  $e$ -gluon fusion in  $ep$  collisions with  $e_8 \rightarrow e g$ . See their Fig. 3 for exclusion plot in the  $m_{e_8}$ - $\Lambda$  plane for  $m_{e_8} = 35$ -220 GeV.

MASS LIMITS for Color Octet Neutrinos ( $\nu_8$ )

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>110	90	<sup>1</sup> BARGER	89 RVUE	$\nu_8$ : $p\bar{p} \rightarrow \nu_8\bar{\nu}_8$

• • • We do not use the following data for averages, fits, limits, etc. • • •  
none 3.8-29.8 95 <sup>2</sup> KIM 90 AMY  $\nu_8$ :  $e^+e^- \rightarrow$  acoplanar jets

none 9-21.9 95 <sup>3</sup> BARTEL 87B JADE  $\nu_8$ :  $e^+e^- \rightarrow$  acoplanar jets

<sup>1</sup> BARGER 89 used ABE 89B limit for events with large missing transverse momentum. Two-body decay  $\nu_8 \rightarrow \nu g$  is assumed.

<sup>2</sup> KIM 90 is at  $E_{cm} = 50$ -60.8 GeV. The same assumptions as in BARTEL 87B are used.

<sup>3</sup> BARTEL 87B is at  $E_{cm} = 46.3$ -46.78 GeV. The limit assumes the  $\nu_8$  pair production cross section to be eight times larger than that of the corresponding heavy neutrino pair production. This assumption is not valid in general for the weak couplings, and the limit can be sensitive to its  $SU(2)_L \times U(1)_Y$  quantum numbers.

MASS LIMITS for  $W_8$  (Color Octet  $W$  Boson)

VALUE (GeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			

<sup>1</sup> ALBAJAR 89 UA1  $p\bar{p} \rightarrow W_8 X, W_8 \rightarrow W g$

<sup>1</sup> ALBAJAR 89 give  $\sigma(W_8 \rightarrow W + \text{jet})/\sigma(W) < 0.019$  (90% CL) for  $m_{W_8} > 220$  GeV.

REFERENCES FOR Searches for Quark and Lepton Compositeness

KHACHATRYAN... 16I	JHEP 1601 166	V. Khachatryan et al.	(CMS Collab.)
AAD 15AP	JHEP 1508 138	G. Aad et al.	(ATLAS Collab.)
AAD 15BY	JHEP 1510 150	G. Aad et al.	(ATLAS Collab.)
AAD 15L	PRL 114 221802	G. Aad et al.	(ATLAS Collab.)
AAD 15V	PR D91 052007	G. Aad et al.	(ATLAS Collab.)
KHACHATRYAN... 15AE	JHEP 1504 025	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN... 15J	PL B746 79	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN... 15V	PR D91 052009	V. Khachatryan et al.	(CMS Collab.)
AAD 14A	PL B728 562	G. Aad et al.	(ATLAS Collab.)
AAD 14AZ	JHEP 1411 104	G. Aad et al.	(ATLAS Collab.)
AAD 14BE	EPL C74 3134	G. Aad et al.	(ATLAS Collab.)
FABBRICHESI 14	PR D89 074028	M. Fabbrichesi, M. Pinamonti, A. Tonerio	(ATLAS Collab.)
KHACHATRYAN... 14	JHEP 1408 173	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN... 14J	PL B738 274	V. Khachatryan et al.	(CMS Collab.)
AAD 13AF	PL B721 171	G. Aad et al.	(ATLAS Collab.)
AAD 13BB	NJP 15 093011	G. Aad et al.	(ATLAS Collab.)
AAD 13D	JHEP 1301 029	G. Aad et al.	(ATLAS Collab.)
AAD 13E	PR D87 015010	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN 13AE	PL B720 309	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AI	PL B722 28	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AJ	PL B723 280	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13AN	PR D87 052017	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN 13K	PR D87 032001	S. Chatrchyan et al.	(CMS Collab.)
AAD 12AB	PL B712 40	G. Aad et al.	(ATLAS Collab.)
AAD 12AZ	PR D85 072003	G. Aad et al.	(ATLAS Collab.)
CHATRCHYAN 12Z	JHEP 1205 055	S. Chatrchyan et al.	(CMS Collab.)
AARON 11C	PL B705 52	F. D. Aaron et al.	(H1 Collab.)
ABAZOV 11F	PRL 106 081801	V. M. Abazov et al.	(DO Collab.)
ABDALLAH 11	EJ C71 1555	J. Abdallah et al.	(DELPHI Collab.)
CHATRCHYAN 11X	PL B704 143	S. Chatrchyan et al.	(CMS Collab.)
AALTONEN 10H	PRL 104 091801	T. Aaltonen et al.	(CDF Collab.)
ABDALLAH 09	EJ C60 1	J. Abdallah et al.	(DELPHI Collab.)
AARON 08	PL B663 382	F. D. Aaron et al.	(H1 Collab.)
SCHAEEL 07A	EJ C49 411	S. Schaeel et al.	(ALEPH Collab.)
ABDALLAH 06C	EJ C45 589	J. Abdallah et al.	(DELPHI Collab.)
ABULENCIA 06L	PRL 96 211801	A. Abulencia et al.	(CDF Collab.)
ABBIENDI 04G	EJ C33 173	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI 04N	PL B602 167	G. Abbiendi et al.	(OPAL Collab.)
ABDALLAH 04N	EJ C37 405	J. Abdallah et al.	(DELPHI Collab.)
ACHARD 03B	PL B568 23	P. Achard et al.	(L3 Collab.)
BABICH 03	EJ C29 103	A. A. Babich et al.	(OPAL Collab.)
ABBIENDI 02G	PL B544 57	G. Abbiendi et al.	(OPAL Collab.)
ACHARD 02D	PL B531 28	P. Achard et al.	(L3 Collab.)
ADLOFF 02	PL B525 9	C. Adloff et al.	(H1 Collab.)
CHEKANOV 02D	PL B549 32	S. Chekanov et al.	(ZEUS Collab.)
AFFOLDER 01I	PRL 87 231803	T. Affolder et al.	(CDF Collab.)
BOURLIKOV 01	PR D64 071701	D. Bourlikov	

CHEUNG 01B	PL B517 167	K. Cheung	
ACCIARRI 00P	PL B489 81	M. Acciarri et al.	(L3 Collab.)
AFFOLDER 00I	PR D62 012004	T. Affolder et al.	(CDF Collab.)
BARATE 98U	EJ C4 571	R. Barate et al.	(ALEPH Collab.)
BARGER 98E	PR D57 391	V. Barger et al.	
MC FARLAND 98	EJ C1 509	K.S. McFarland et al.	(CCFR/NuTeV Collab.)
DIAZCRUZ 94	PR D49 R2149	J.L. Diaz Cruz, O.A. Sampayo	(CINUV)
ABT 93	NP B396 3	I. Abt et al.	(H1 Collab.)
ADRIANI 93M	PRPL 236 1	O. Adriani et al.	(L3 Collab.)
BARDADIN... 92	ZPHY C55 163	M. Bardadin-Otwinowska	(CLER)
DECAMP 92	PRPL 216 253	D. Decamp et al.	(ALEPH Collab.)
PDG 92	PR D45 51	K. Hikasa et al.	(KEK, LBL, BOST+)
ABREU 91F	NP B367 511	P. Abreu et al.	(DELPHI Collab.)
KIM 90	PL B240 243	G.N. Kim et al.	(AMY Collab.)
ABE 89B	PRL 62 1825	F. Abe et al.	(CDF Collab.)
ABE 89D	PRL 63 1447	F. Abe et al.	(CDF Collab.)
ABE 89J	ZPHY C45 175	K. Abe et al.	(VENUS Collab.)
ALBAJAR 89	ZPHY C44 15	C. Alajar et al.	(UA1 Collab.)
BARGER 89	PL B220 464	V. Barger et al.	(WISC, KEK)
DORENBOS... 89	ZPHY C41 567	J. Dorenbosch et al.	(CHARM Collab.)
BARTEL 87B	ZPHY C36 15	W. Bartel et al.	(JADE Collab.)
GRIFOLS 86	PL B68B 264	J.A. Grifols, S. Peris	(BARC)
JODIDIO 86	PR D34 1967	A. Jodidio et al.	(LBL, NWES, TRIU)
Also	PR D37 237 (erratum)	A. Jodidio et al.	(LBL, NWES, TRIU)
RENARD 82	PL B116B 264	F.M. Renard	(CERN)

Extra Dimensions

For explanation of terms used and discussion of significant model dependence of following limits, see the "Extra Dimensions" review. Footnotes describe originally quoted limit.  $\delta$  indicates the number of extra dimensions.

Limits not encoded here are summarized in the "Extra Dimensions" review, where the latest unpublished results are also described.

EXTRA DIMENSIONS

Updated September 2015 by John Parsons (Columbia University) and Alex Pomarol (Universitat Autònoma de Barcelona)

I Introduction

Proposals for a spacetime with more than three spatial dimensions date back to the 1920's, mainly through the work of Kaluza and Klein, in an attempt to unify the forces of nature [1]. Although their initial idea failed, the formalism that they and others developed is still useful nowadays. Around 1980, string theory proposed again to enlarge the number of space dimensions, this time as a requirement for describing a consistent theory of quantum gravity. The extra dimensions were supposed to be compactified at a scale close to the Planck scale, and thus not testable experimentally in the near future.

A different approach was given by Arkani-Hamed, Dimopoulos and Dvali (ADD) in their seminal paper in 1998 [2], where they showed that the weakness of gravity could be explained by postulating two or more extra dimensions in which only gravity could propagate. The size of these extra dimensions should range between roughly a millimeter and  $\sim 1/\text{TeV}$ , leading to possible observable consequences in current and future experiments. A year later, Randall and Sundrum (RS) [3] found a new possibility using a warped geometry, postulating a five-dimensional Anti-de Sitter (AdS) spacetime with a compactification scale of order TeV. The origin of the smallness of the electroweak scale versus the Planck scale was explained by the gravitational redshift factor present in the warped AdS metric. As in the ADD model, originally only gravity was assumed to propagate in the extra dimensions, although it was soon clear that this was not necessary in warped extra-dimensions and also the SM gauge fields [4] and SM fermions [5,6] could propagate in the five-dimensional space.

# Searches Particle Listings

## Extra Dimensions

The physics of warped extra-dimensional models has an alternative interpretation by means of the AdS/CFT correspondence [7]. Models with warped extra dimensions are related to four-dimensional strongly-interacting theories, allowing an understanding of the properties of five-dimensional fields as those of four-dimensional composite states [8]. This approach has opened new directions for tackling outstanding questions in particle physics, such as the flavor problem, grand unification, and the origin of electroweak symmetry breaking or supersymmetry breaking.

### Experimental

**Constraints:** Constraints on extra-dimensional models arise from astrophysical and cosmological considerations. In addition, as we will show below, tabletop experiments exploring gravity at sub-mm distances restrict certain models. Collider limits on extra-dimensional models are dominated by Run 1 LHC results. This review includes the most recent limits, most of which are published results based on LHC data collected in 2012 at a center-of-mass energy of 8 TeV. In addition, there are a few preliminary 8 TeV results, which can be found on the WWW pages of public ATLAS [9] and CMS [10]. Unless otherwise stated, all LHC results use the full  $\sim 20 \text{ fb}^{-1}$  samples of 8 TeV collisions recorded in 2012. Run 2 LHC results, with operations at 13 TeV, will greatly extend the sensitivity of the collider searches for evidence of extra dimensions, particularly once comparable integrated luminosities (and eventually much more) have been accumulated. For some models, much smaller data samples already suffice; indeed a few preliminary 13 TeV analyses focusing on strong gravity signatures have been presented at the LHCP Conference in early September 2015, and are included here since they already surpass the sensitivity of the 8 TeV results, despite using very early Run 2 data samples of less than  $0.1 \text{ fb}^{-1}$ .

**Kaluza-Klein Theories:** Field theories with compact extra dimensions can be written as theories in ordinary four dimensions (4D) by performing a Kaluza-Klein (KK) reduction. As an illustration, consider a simple example, namely a field theory of a complex scalar in flat five-dimensional (5D) spacetime. The action will be given by <sup>†</sup>

$$S_5 = - \int d^4x dy M_5 \left[ |\partial_\mu \phi|^2 + |\partial_y \phi|^2 + \lambda_5 |\phi|^4 \right], \quad (1)$$

where  $y$  refers to the extra (fifth) dimension. A universal scale  $M_5$  has been extracted in front of the action in order to keep the 5D field with the same mass-dimension as in 4D. This theory is perturbative for energies  $E \lesssim \ell_5 M_5 / \lambda_5$  where  $\ell_5 = 24\pi^3$  [11].

Let us now consider that the fifth dimension is compact with the topology of a circle  $S^1$  of radius  $R$ , which corresponds to the identification of  $y$  with  $y + 2\pi R$ . In such a case, the 5D complex scalar field can be expanded in a Fourier series:

$$\phi(x, y) = \frac{1}{\sqrt{2\pi R M_5}} \sum_{n=-\infty}^{\infty} e^{iny/R} \phi^{(n)}(x),$$

<sup>†</sup> Our convention for the metric is  $\eta_{MN} = \text{Diag}(-1, 1, 1, 1, 1)$ .

that, inserted in Eq. (1) and integrating over  $y$ , gives

$$S_5 = S_4^{(0)} + S_4^{(n)},$$

where

$$S_4^{(0)} = - \int d^4x \left[ |\partial_\mu \phi^{(0)}|^2 + \lambda_4 |\phi^{(0)}|^4 \right], \text{ and} \quad (2)$$

$$S_4^{(n)} = - \int d^4x \sum_{n \neq 0} \left[ |\partial_\mu \phi^{(n)}|^2 + \left(\frac{n}{R}\right)^2 |\phi^{(n)}|^2 \right] + \text{quartic int.}$$

The  $n = 0$  mode self-coupling is given by

$$\lambda_4 = \frac{\lambda_5}{2\pi R M_5}. \quad (3)$$

The above action corresponds to a 4D theory with a massless scalar  $\phi^{(0)}$ , referred to as the zero-mode, and an infinite tower of massive modes  $\phi^{(n)}$ , known as KK modes. The KK reduction thus allows a treatment of 5D theories as 4D field theories with an infinite number of fields. At energies smaller than  $1/R$ , the KK modes can be neglected, leaving the zero-mode action of Eq. (2). The strength of the interaction of the zero-mode, given by Eq. (3), decreases as  $R$  increases. Thus, for a large extra dimension  $R \gg 1/M_5$ , the massless scalar is weakly coupled.

## II Large Extra Dimensions for Gravity

### II.1 The ADD Scenario

The ADD scenario [2,12,13] assumes a  $D = 4 + \delta$  dimensional spacetime, with  $\delta$  compactified spatial dimensions. The weakness of gravity arises since it propagates in the higher-dimensional space. The SM is assumed to be localized in a 4D subspace, a 3-brane, as can be found in certain string constructions [14]. Gravity is described by the Einstein-Hilbert action in  $D = 4 + \delta$  spacetime dimensions

$$S_D = - \frac{\bar{M}_D^{2+\delta}}{2} \int d^4x d^\delta y \sqrt{-g} \mathcal{R} + \int d^4x \sqrt{-g_{\text{ind}}} \mathcal{L}_{\text{SM}}, \quad (4)$$

where  $x$  labels the ordinary four coordinates,  $y$  the  $\delta$  extra coordinates,  $g$  refers to the determinant of the  $D$ -dimensional metric whose Ricci scalar is defined by  $\mathcal{R}$ , and  $\bar{M}_D$  is the reduced Planck scale of the  $D$ -dimensional theory. In the second term of Eq. (4), which gives the gravitational interactions of SM fields, the  $D$ -dimensional metric reduces to the induced metric on the 3-brane where the SM fields propagate. The extra dimensions are assumed to be flat and compactified in a volume  $V_\delta$ . As an example, consider a toroidal compactification of equal radii  $R$  and volume  $V_\delta = (2\pi R)^\delta$ . After a KK reduction, one finds that the fields that couple to the SM are the spin-2 gravitational field  $G_{\mu\nu}(x, y)$  and a tower of spin-1 KK graviscalars [15]. The graviscalars, however, only couple to SM fields through the trace of the energy-momentum tensor, resulting in weaker couplings to the SM fields. The Fourier expansion of the spin-2 field is given by

$$G_{\mu\nu}(x, y) = G_{\mu\nu}^{(0)}(x) + \frac{1}{\sqrt{V_\delta}} \sum_{\vec{n} \neq 0} e^{i\vec{n} \cdot \vec{y}/R} G_{\mu\nu}^{(\vec{n})}(x), \quad (5)$$

where  $\vec{y} = (y_1, y_2, \dots, y_\delta)$  are the extra-dimensional coordinates and  $\vec{n} = (n_1, n_2, \dots, n_\delta)$ . Eq. (5) contains a massless state, the 4D graviton, and its KK tower with masses  $m_n^2 = |\vec{n}|^2/R^2$ . At energies below  $1/R$  the action is that of the zero-mode

$$S_4^{(0)} = -\frac{\bar{M}_D^{2+\delta}}{2} \int d^4x V_\delta \sqrt{-g^{(0)}} \mathcal{R}^{(0)} + \int d^4x \sqrt{-g_{\text{ind}}^{(0)}} \mathcal{L}_{\text{SM}},$$

where we can identify the 4D reduced Planck mass,  $M_P \equiv G_N/\sqrt{8\pi} \simeq 2.4 \times 10^{18}$  GeV, as a function of the  $D$ -dimensional parameters:

$$M_P^2 = V^\delta \bar{M}_D^{2+\delta} \equiv R^\delta M_D^{2+\delta}. \quad (6)$$

Fixing  $M_D$  at around the electroweak scale  $M_D \sim \text{TeV}$  to avoid introducing a new mass-scale in the model, Eq. (6) gives a prediction for  $R$ :

$$\delta = 1, 2, \dots, 6 \rightarrow R \sim 10^9 \text{ km}, 0.5 \text{ mm}, \dots, 0.1 \text{ MeV}^{-1}. \quad (7)$$

The option  $\delta = 1$  is clearly ruled out, as it leads to modifications of Newton's law at solar system distances. However this is not the case for  $\delta \geq 2$ , and possible observable consequences can be sought in present and future experiments.

Consistency of the model requires a stabilization mechanism for the radii of the extra dimensions, to the values shown in Eq. (7). The fact that we need  $R \gg 1/M_D$  leads to a new hierarchy problem, the solution of which might require imposing supersymmetry in the extra-dimensional bulk [16].

## II.2 Tests of the Gravitational Force Law at Sub-mm Distances

The KK modes of the graviton give rise to deviations from Newton's law of gravitation for distances  $\lesssim R$ . Such deviations are usually parametrized by a modified Newtonian potential of the form

$$V(r) = -G_N \frac{m_1 m_2}{r} \left[ 1 + \alpha e^{-r/\lambda} \right]. \quad (8)$$

For a 2-torus compactification,  $\alpha = 16/3$  and  $\lambda = R$ . Searches for deviations from Newton's law of gravitation have been performed in several experiments. Ref. [17] gives the present constraints:  $R < 37 \mu\text{m}$  at 95% CL for  $\delta = 2$ , corresponding to  $M_D > 3.6 \text{ TeV}$ .

## II.3 Astrophysical and Cosmological Constraints

The light KK gravitons could be copiously produced in stars, carrying away energy. Ensuring that the graviton luminosity is low enough to preserve the agreement of stellar models with observations provides powerful bounds on the scale  $M_D$ . The most stringent arises from supernova SN1987A, giving  $M_D > 27$  (2.4) TeV for  $\delta = 2$  (3) [18]. After a supernova explosion, most of the KK gravitons stay gravitationally trapped in the remnant neutron star. The requirement that neutron stars are not excessively heated by KK decays into photons leads to  $M_D > 1700$  (76) TeV for  $\delta = 2$  (3) [19].

Cosmological constraints are also quite stringent [20]. To avoid overclosure of the universe by relic gravitons one needs  $M_D > 7 \text{ TeV}$  for  $\delta = 2$ . Relic KK gravitons decaying into photons contribute to the cosmic diffuse gamma radiation, from which one can derive the bound  $M_D > 100 \text{ TeV}$  for  $\delta = 2$ .

We must mention however that bounds coming from the decays of KK gravitons into photons can be reduced if we assume that KK gravitons decay mainly into other non-SM states. This could happen, for example, if there were other 3-branes with hidden sectors residing on them [12].

## II.4 Collider Signals

### II.4a Graviton and Other Particle Production

Although each KK graviton has a purely gravitational coupling, suppressed by  $1/M_P$ , inclusive processes in which one sums over the almost continuous spectrum of available gravitons have cross sections suppressed only by powers of  $M_D$ . Processes involving gravitons are therefore detectable in collider experiments if  $M_D \sim \text{TeV}$ . A number of experimental searches for evidence of large extra dimensions have been performed at colliders, and interpreted in the context of the ADD model.

One signature arises from direct graviton emission. By making a derivative expansion of Einstein gravity, one can construct an effective theory, valid for energies much lower than  $M_D$ , and use it to make predictions for graviton-emission processes at colliders [15,21,22]. Gravitons produced in the final state would escape detection, giving rise to missing transverse energy ( $\cancel{E}_T$ ). The results quoted below are 95% CL lower limits on  $M_D$  for a range of values of  $\delta$  between 2 and 6, with more stringent limits corresponding to lower  $\delta$  values.

At hadron colliders, experimentally sensitive channels include the jet ( $j$ ) +  $\cancel{E}_T$  and  $\gamma$  +  $\cancel{E}_T$  final states. The CMS analysis of the  $j$  +  $\cancel{E}_T$  final state assumes k-factors varying between 1.4 for  $\delta = 2, 3$  and down to 1.2 for  $\delta = 6$  to account for next-to-leading order (NLO) contributions to the signal cross sections, and sets limits of  $M_D > 3.53 - 6.09 \text{ TeV}$  [23]. ATLAS  $j$  +  $\cancel{E}_T$  results provide limits of  $M_D > 3.06 - 5.25 \text{ TeV}$  [24], assuming leading order (LO) cross sections. For these analyses, the LHC experiments handle somewhat differently the issue that the effective theory is only valid for energies much less than  $M_D$ : the ATLAS results are quoted for the full space, and include the information that suppressing the graviton cross section by a factor  $M_D^4/\hat{s}^2$  for  $\sqrt{\hat{s}} > M_D$ , where  $\sqrt{\hat{s}}$  is the parton-level center-of-mass energy of the hard collision, weakens the limits on  $M_D$  by a negligible amount ( $\sim 3\%$ ) for  $\delta = 2$  ( $\delta = 6$ ). CMS considers the impact of simply truncating the differential cross section to remove the contribution from events where  $\sqrt{\hat{s}} > M_D$ , and shows that the effect of the truncation changes the cross section by a maximum of 11%. Less stringent limits are obtained by both ATLAS [25] and CMS [26] from analyses of the  $\gamma$  +  $\cancel{E}_T$  final state.

In models in which the ADD scenario is embedded in a string theory at the TeV scale [14], we expect the string scale  $M_s$  to be smaller than  $M_D$ , and therefore expect production of string resonances at the LHC [27]. A preliminary CMS Run 2 result from analysing the dijet invariant mass distribution for  $42 \text{ pb}^{-1}$  of 13 TeV data excludes string resonances that decay predominantly to  $q+g$  with masses below 5.1 TeV [28], already slightly extending their lower limit [29] of 5.0 TeV obtained

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using their complete 20 fb<sup>-1</sup> sample of 8 TeV collisions. ATLAS dijet analyses include a published result using their full 8 TeV dataset [30] and a preliminary Run 2 result using 80 pb<sup>-1</sup> of 13 TeV data [31], and provide their results in the context of model-independent limits on the cross section times acceptance for generic resonances of a variety of possible widths.

#### II.4b Virtual graviton effects

One can also search for virtual graviton effects, the calculation of which however depends on the ultraviolet cut-off of the theory and is therefore very model dependent. In the literature, several different formulations exist [15,22,32] for the dimension-eight operator for gravity exchange at tree level:

$$\mathcal{L}_8 = \pm \frac{4}{M_{TT}^4} \left( T_{\mu\nu} T^{\mu\nu} - \frac{1}{\delta + 2} T_\mu^\mu T_\nu^\nu \right), \quad (9)$$

where  $T_{\mu\nu}$  is the energy-momentum tensor and  $M_{TT}$  is related to  $M_D$  by some model-dependent coefficient [33]. The relations with the parametrizations of Refs. [32] and [15] are, respectively,  $M_{TT} = M_S$  and  $M_{TT} = (2/\pi)^{1/4} \Lambda_T$ . The experimental results below are given as 95% CL lower limits on  $M_{TT}$ , including in some cases the possibility of both constructive or destructive interference, depending on the sign chosen in Eq. (9).

The most stringent limits arise from LHC analyses of the dijet angular distribution. Using their full 8 TeV dataset, CMS [34] obtains results that correspond to an approximate limit of  $M_{TT} > 6.3$  TeV. An ATLAS analysis [35] of 17 fb<sup>-1</sup> of 8 TeV collisions would provide similar sensitivity, but quotes the results only in the context of limits on contact interactions. The next most restrictive results are obtained by LHC analyses combining the dielectron and dimuon final states, with both experiments providing similar limits of approximately  $M_{TT} > 3.7$  TeV. The ATLAS [36] (CMS [37]) dilepton results assume LO (NLO) signal cross section values.

At the one-loop level, gravitons can also generate dimension-six operators with coefficients that are also model dependent. Experimental bounds on these operators can also give stringent constraints on  $M_D$  [33].

#### II.4c Black Hole Production

The physics at energies  $\sqrt{s} \sim M_D$  is sensitive to the details of the unknown quantum theory of gravity. Nevertheless, in the transplanckian regime,  $\sqrt{s} \gg M_D$ , one can rely on a semiclassical description of gravity to obtain predictions. An interesting feature of transplanckian physics is the creation of black holes [38]. A black hole is expected to be formed in a collision in which the impact parameter is smaller than the Schwarzschild radius [39]:

$$R_S = \frac{1}{M_D} \left[ \frac{2^\delta \pi^{\delta-3/2}}{\delta + 2} \Gamma\left(\frac{\delta + 3}{2}\right) \frac{M_{BH}}{M_D} \right]^{1/(\delta+1)}, \quad (10)$$

where  $M_{BH}$  is the mass of the black hole, which would roughly correspond to the total energy in the collision. The cross section for black hole production can be estimated to be of the same order as the geometric area  $\sigma \sim \pi R_S^2$ . For  $M_D \sim \text{TeV}$ , this gives

a production of  $\sim 10^7$  black holes at the  $\sqrt{s} = 14$  TeV LHC with an integrated luminosity of 30 fb<sup>-1</sup> [38]. A black hole would provide a striking experimental signature since it is expected to thermally radiate with a Hawking temperature  $T_H = (\delta + 1)/(4\pi R_S)$ , and therefore would evaporate democratically into all SM states. Nevertheless, given the present constraints on  $M_D$ , the LHC will not be able to reach energies much above  $M_D$ . This implies that predictions based on the semiclassical approximation could receive sizable modifications from model-dependent quantum-gravity effects.

The most stringent limits on microscopic black holes arise from LHC searches which observed no excesses above the SM background in high-multiplicity final states. The results are usually quoted as model-independent limits on the cross section for new physics in the final state and kinematic region analyzed. These results can then be used to provide constraints of models of low-scale gravity and weakly-coupled string theory. In addition, limits are sometimes quoted on particular implementations of models, which are used as benchmarks to illustrate the sensitivity. A preliminary Run 2 ATLAS search [40] for an excess of events with multiple high transverse momentum objects, including charged leptons and jets, using 80 pb<sup>-1</sup> of 13 TeV data, excludes semiclassical black holes below masses of  $\sim 7.3$  TeV for  $M_D = 2$  TeV and  $\delta = 6$ , extending by almost 1.5 TeV the limit of the corresponding analysis [41] applied to the full 20 fb<sup>-1</sup> 8 TeV dataset. Another preliminary Run 2 ATLAS analysis [42], again using 80 pb<sup>-1</sup> of 13 TeV data, looks at very high transverse energy multijet events and excludes black hole masses in the range 7.5 – 8.5 TeV, depending on  $M_D$ , for  $\delta = 6$ , again extending the limits of the corresponding 8 TeV analysis [43] by 0.5-1.5 TeV. The 8 TeV ATLAS analysis [44] of the track multiplicity in same-sign dimuon events provides lower mass limits of 5.1 - 5.7 TeV for  $M_D = 1.5$  TeV, with the range of limits depending on details of the model and also the number of extra dimensions. A CMS analysis [45] of multi-object final states using 12 fb<sup>-1</sup> of 8 TeV data provides similar limits, extending out to values of  $M_D \sim 5$  TeV.

For black hole masses near  $M_D$ , the semi-classical approximation is not valid, and one instead expects quantum black holes (QBH) that decay primarily into two-body final states [46]. LHC Run 1 results at 8 TeV provide lower limits on quantum black hole masses of order 4.6 - 6.3 TeV, depending on the details of the model. Searches that consider interpretations in terms of QBH limits include the CMS multi-object [45] analysis, as well as their dijet analysis [29]. ATLAS results include, in addition to their dijet analysis [30], searches in the photon+jet [47] and lepton+jet [48] final states. The preliminary Run 2 ATLAS dijet analysis [31] using 80 pb<sup>-1</sup> of 13 TeV collisions extends the exclusions to 6.5 – 6.8 TeV, depending on the model details.

In weakly-coupled string models the semiclassical description of gravity fails in the energy range between  $M_s$  and  $M_s/g_s^2$  where stringy effects are important. In this regime one expects, instead of black holes, the formation of string balls, made of



highly excited long strings, that could be copiously produced at the LHC for  $M_s \sim \text{TeV}$  [49], and would evaporate thermally at the Hagedorn temperature giving rise to high-multiplicity events. The same analyses used to search for black holes can be interpreted in the context of string balls. For example, for the case of  $\delta = 6$  and with model parameters fixed to values of  $g_s = 0.4$ ,  $M_D = 1.5 \text{ TeV}$ , and  $M_s = M_D/1.26 = 1.2 \text{ TeV}$ , the ATLAS same-sign dimuon [44] (multiple high transverse momentum object [41]) analysis excludes string balls with minimal masses below 5.3 (5.7) TeV. The CMS multi-object [45] analysis excludes the production of string balls with a minimum mass below  $\sim 5.5 \text{ TeV}$  for  $g_s = 0.4$ ,  $M_D$  in the range of 1.4–2.1 TeV, and  $M_s = M_D/1.25$ .

### III TeV-Scale Extra Dimensions

#### III.1 Warped Extra Dimensions

The RS model [3] is the most attractive setup of warped extra dimensions at the TeV scale, since it provides an alternative solution to the hierarchy problem. The RS model is based on a 5D theory with the extra dimension compactified in an orbifold,  $S^1/Z_2$ , a circle  $S^1$  with the extra identification of  $y$  with  $-y$ . This corresponds to the segment  $y \in [0, \pi R]$ , a manifold with boundaries at  $y = 0$  and  $y = \pi R$ . Let us now assume that this 5D theory has a cosmological constant in the bulk  $\Lambda$ , and on the two boundaries  $\Lambda_0$  and  $\Lambda_{\pi R}$ :

$$S_5 = - \int d^4x dy \left\{ \sqrt{-g} \left[ \frac{1}{2} M_5^3 \mathcal{R} + \Lambda \right] + \sqrt{-g_0} \delta(y) \Lambda_0 + \sqrt{-g_{\pi R}} \delta(y - \pi R) \Lambda_{\pi R} \right\}, \quad (11)$$

where  $g_0$  and  $g_{\pi R}$  are the values of the determinant of the induced metric on the two respective boundaries. Einstein's equations can be solved, giving in this case the metric

$$ds^2 = a(y)^2 dx^\mu dx^\nu \eta_{\mu\nu} + dy^2, \quad a(y) = e^{-ky}, \quad (12)$$

where  $k = \sqrt{-\Lambda/6M_5^3}$ . Consistency of the solution requires  $\Lambda_0 = -\Lambda_{\pi R} = -\Lambda/k$ . The metric in Eq. (12) corresponds to a 5D AdS space. The factor  $a(y)$  is called the “warp” factor and determines how 4D scales change as a function of the position in the extra dimension. In particular, this implies that energy scales for 4D fields localized at the boundary at  $y = \pi R$  are red-shifted by a factor  $e^{-k\pi R}$  with respect to those localized at  $y = 0$ . For this reason, the boundaries at  $y = 0$  and  $y = \pi R$  are usually referred to as the ultraviolet (UV) and infrared (IR) boundaries, respectively.

As in the ADD case, we can perform a KK reduction and obtain the low-energy effective theory of the 4D massless graviton. In this case we obtain

$$M_P^2 = \int_0^{\pi R} dy e^{-2ky} M_5^3 = \frac{M_5^3}{2k} (1 - e^{-2k\pi R}). \quad (13)$$

Taking  $M_5 \sim k \sim M_P$ , we can generate an IR-boundary scale of order  $ke^{-k\pi R} \sim \text{TeV}$  for an extra dimension of radius  $R \simeq 11/k$ . Mechanisms to stabilize  $R$  to this value have been proposed [50]

that, contrary to the ADD case, do not require introducing any new small or large parameter. Therefore a natural solution to the hierarchy problem can be achieved in this framework if the Higgs field, whose vacuum expectation value (VEV) is responsible for electroweak symmetry breaking, is localized at the IR-boundary where the effective mass scales are of order TeV. The radion field is generically heavy in models with a stabilized  $R$ . Nevertheless, it has been recently discussed that under some conditions a naturally light radion can arise [51]. In these cases the radion is identified with the dilaton, the Goldstone boson associated to the spontaneous breaking of scale invariance, and its mass can be naturally below  $ke^{-k\pi R} \sim \text{TeV}$ .

In the RS model [3], all the SM fields were assumed to be localized on the IR-boundary. Nevertheless, for the hierarchy problem, only the Higgs field has to be localized there. SM gauge bosons and fermions can propagate in the 5D bulk [4,5,6,52]. By performing a KK reduction from the 5D action of a gauge boson, we find [4]

$$\frac{1}{g_4^2} = \int_0^{\pi R} dy \frac{1}{g_5^2} = \frac{\pi R}{g_5^2},$$

where  $g_D$  ( $D = 4, 5$ ) is the gauge coupling in  $D$ -dimensions. Therefore the 4D gauge couplings can be of order one, as is the case of the SM, if one demands  $g_5^2 \sim \pi R$ . Using  $kR \sim 10$  and  $g_4 \sim 0.5$ , one obtains the 5D gauge coupling

$$g_5 \sim 4/\sqrt{k}. \quad (14)$$

Boundary kinetic terms for the gauge bosons can modify this relation, allowing for larger values of  $g_5\sqrt{k}$ .

Fermions propagating in a warped extra dimension have 4D massless zero-modes with wavefunctions which vary as  $f_0 \sim \exp[(1/2 - c_f)ky]$ , where  $c_fk$  is their 5D mass [53,6]. Depending on the free parameter  $c_fk$ , fermions can be localized either towards the UV-boundary ( $c_f > 1/2$ ) or IR-boundary ( $c_f < 1/2$ ). Since the Higgs boson is localized on the IR-boundary, one can generate exponentially suppressed Yukawa couplings by having the fermion zero-modes localized towards the UV-boundary, generating naturally the light SM fermion spectrum [6]. A large overlap with the wavefunction of the Higgs is needed for the top quark, in order to generate its large mass, thus requiring it to be localized towards the IR-boundary. In conclusion, the large mass hierarchies present in the SM fermion spectrum can be easily obtained in warped models via suitable choices of the order-one parameters  $c_f$  [54]. In these scenarios, deviations in flavor physics from the SM predictions are expected to arise from flavor-changing KK gluon couplings [55], putting certain constraints on the parameters of the models and predicting new physics effects to be observed in  $B$ -physics processes [56].

The masses of the KK states can also be calculated. One finds [6]

$$m_n \simeq \left( n + \frac{\alpha}{2} - \frac{1}{4} \right) \pi k e^{-\pi k R}, \quad (15)$$

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where  $n = 1, 2, \dots$  and  $\alpha = \{|c_f - 1/2|, 0, 1\}$  for KK fermions, KK gauge bosons and KK gravitons, respectively. Their masses are of order  $ke^{-\pi kR} \sim \text{TeV}$ ; the first KK state of the gauge bosons would be the lightest, while gravitons are expected to be the heaviest.

### III.1a Models of Electroweak Symmetry Breaking

Theories in warped extra dimensions can be used to implement symmetry breaking at low energies by boundary conditions [57]. For example, for a  $U(1)$  gauge symmetry in the 5D bulk, this can be easily achieved by imposing a Dirichlet boundary condition on the IR-boundary for the gauge-boson field,  $A_\mu|_{y=\pi R} = 0$ . This makes the zero-mode gauge boson get a mass, given by  $m_A = g_4 \sqrt{2k/g_5^2} e^{-\pi kR}$ . A very different situation occurs if the Dirichlet boundary condition is imposed on the UV-boundary,  $A_\mu|_{y=0} = 0$ . In this case the zero-mode gauge boson disappears from the spectrum. Finally, if a Dirichlet boundary condition is imposed on the two boundaries, one obtains a massless 4D scalar corresponding to the fifth component of the 5D gauge boson,  $A_5$ . Thus, different scenarios can be implemented by appropriately choosing the 5D bulk gauge symmetry,  $\mathcal{G}_5$ , and the symmetries to which it reduces on the UV and IR-boundary,  $\mathcal{H}_{UV}$  and  $\mathcal{H}_{IR}$ , respectively. In all cases the KK spectrum comes in representations of the group  $\mathcal{G}_5$ .

The recent discovery of a light Higgs boson with  $m_H \sim 125 \text{ GeV}$  [58] rules out Higgsless 5D models for electroweak symmetry breaking [59]. This discovery, however, is consistent with 5D composite Higgs models where a light Higgs boson is present in the spectrum.

**Composite Higgs models:** Warped extra dimensions can give rise to scenarios, often called gauge-Higgs unified models, where the Higgs boson appears as the fifth component of a 5D gauge boson,  $A_5$ . The Higgs mass is protected by the 5D gauge invariance and can only get a nonzero value from non-local one-loop effects [60]. To guarantee the relation  $M_W^2 \simeq M_Z^2 \cos^2 \theta_W$ , a custodial  $SU(2)_V$  symmetry is needed in the bulk and IR-boundary [61]. The simplest realization [62] has

$$\begin{aligned}\mathcal{G}_5 &= SU(3)_c \times SO(5) \times U(1)_X, \\ \mathcal{H}_{IR} &= SU(3)_c \times SO(4) \times U(1)_X, \\ \mathcal{H}_{UV} &= G_{SM}.\end{aligned}$$

The Higgs boson gets a potential at the one-loop level that triggers a VEV, breaking the electroweak symmetry. In these models there is a light Higgs boson whose mass can be around 125 GeV, as required by the recently discovered Higgs boson [58]. This state, as will be explained in Sec. III.2, behaves as a composite pseudo-Goldstone boson with couplings that deviate from the SM Higgs [63]. The present experimental determination of the Higgs couplings at the LHC, that agrees with the SM predictions, put important constraints on these scenarios [58]. The lightest KK modes of the model are color fermions with charges  $Q = -1/3, 2/3$  and  $5/3$  [64].

### III.1b Constraints from Electroweak Precision Tests

Models in which the SM gauge bosons propagate in 1/TeV-sized extra dimensions give generically large corrections to electroweak observables. When the SM fermions are confined on a boundary these corrections are universal and can be parametrized by four quantities:  $\hat{S}$ ,  $\hat{T}$ ,  $W$  and  $Y$ , as defined in Ref. [65]. For warped models, where the 5D gauge coupling of Eq. (14) is large, the most relevant parameter is  $\hat{T}$ , which gives the bound  $m_{KK} \gtrsim 10 \text{ TeV}$  [52]. When a custodial symmetry is imposed [61], the main constraint comes from the  $\hat{S}$  parameter, requiring  $m_{KK} \gtrsim 3 \text{ TeV}$ , independent of the value of  $g_5$ . Corrections to the  $Zb_L\bar{b}_L$  coupling can also be important [52], especially in warped models for electroweak symmetry breaking as the ones described above.

### III.1c Kaluza-Klein Searches

The main prediction of 1/TeV-sized extra dimensions is the presence of a discretized KK spectrum, with masses around the TeV scale, associated with the SM fields that propagate in the extra dimension.

In the RS model [3], only gravity propagates in the 5D bulk. Experimental searches have been performed for the lightest KK graviton through its decay to a variety of SM particle-antiparticle pairs. The results are usually interpreted in the plane of the dimensionless coupling  $k/M_P$  versus  $m_1$ , where  $M_P$  is the reduced Planck mass defined previously and  $m_1$  is the mass of the lightest KK excitation of the graviton. Since the AdS curvature  $\sim k$  cannot exceed the cut-off scale of the model, which is estimated to be  $\ell_5^{1/3} M_5$  [33], one must demand  $k \ll \sqrt{2\ell_5} M_P$ . The results quoted below are 95% CL lower limits on the KK graviton mass for a coupling  $k/M_P = 0.1$ .

The most stringent limits currently arise from LHC searches for resonances in the dilepton and diphoton final states, using the full samples of 8 TeV collisions. The CMS [66] and ATLAS [67] dilepton analyses, combining results from the  $ee$  and  $\mu\mu$  channels, exclude gravitons with masses below 2.73 TeV and 2.68 TeV, respectively. Similar results are obtained in the  $\gamma\gamma$  final state, which is quite powerful since it has a branching fraction twice that of any individual lepton flavor. The ATLAS  $\gamma\gamma$  analysis [68] provides a lower limit on the graviton mass of 2.66 TeV, while a preliminary CMS result [69] excludes gravitons below 2.78 TeV. Less stringent limits on the KK graviton mass come from 8 TeV analyses of the dijet [29],  $HH$  [70] and  $VV$  [71] final states, where  $V$  can represent either a  $W$  or  $Z$  boson. Experimental searches for the radion [70], through its production via gluon fusion and decaying to  $HH$ , exclude masses from 300 to 1100 GeV for a decay constant of 1 TeV.

In warped extra-dimensional models in which the SM fields propagate in the 5D bulk, the couplings of the KK graviton to  $ee/\mu\mu/\gamma\gamma$  are suppressed [72], and the above bounds do not apply. Furthermore, the KK graviton is the heaviest KK state (see Eq. (15)), and therefore experimental searches for KK gauge bosons and fermions are more appropriate discovery channels in these scenarios. For the scenarios discussed above in which only the Higgs boson and the top quark are localized

close to the IR-boundary, the KK gauge bosons mainly decay into top quarks, longitudinal  $W/Z$  bosons, and Higgs bosons. Couplings to light SM fermions are suppressed by a factor  $g/\sqrt{g_5^2 k} \sim 0.2$  [6] for the value of Eq. (14) that is considered from now on. Searches have been made for evidence of the lightest KK excitation of the gluon, through its decay to  $t\bar{t}$  pairs. The searches take into account the natural KK gluon width, which is typically  $\sim 15\%$  of its mass. The decay of a heavy particle to  $t\bar{t}$  would tend to produce highly boosted top (anti-)quarks in the final state. Products of the subsequent top decays would therefore tend to be close to each other in the detector. In the case of  $t \rightarrow Wb \rightarrow jjb$  decays, the three jets could overlap one another and not be individually reconstructed with the standard jet algorithms, while  $t \rightarrow Wb \rightarrow \ell\nu b$  decays could result in the lepton failing standard isolation requirements due to its proximity to the  $b$ -jet; in both cases, the efficiency for properly reconstructing the final state would fall as the mass of the original particle increases. To avoid the loss in sensitivity which would result, a number of techniques, known generally as “top quark tagging”, have been developed to reconstruct and identify highly boosted top quarks, for example by using a single “wide” jet to contain all the decay products of a hadronic top decay. The large backgrounds from QCD jets can then be reduced by requiring the “jet mass” be consistent with that of a top quark, and also by examining the substructure of the wide jet for indication that it resulted from the hadronic decay of a top quark. These techniques are key to extending to very high masses the range of accessible resonances decaying to  $t\bar{t}$  pairs. The CMS analysis [73] combines results from the dilepton, lepton-plus-jets and fully hadronic final states and excludes KK gluons with masses below 2.8 TeV. An ATLAS analysis [74] of the lepton-plus-jets final state excludes KK gluon masses below 2.2 TeV. The results are not directly comparable between the two LHC experiments, since they employ in their respective analyses different implementations of the theoretical model.

A gauge boson KK excitation could be also sought through its decay to longitudinal  $W/Z$  bosons. While searches for  $WZ$  resonances have been used to set limits on sequential SM  $W'$  bosons [75] or other models, as yet no  $WZ$  experimental results have been interpreted in the context of warped extra dimensions. The decay to a pair of intermediate vector bosons has, however, been exploited to search for KK gravitons in models in which the SM fields propagate in the 5D bulk. The analyses typically reconstruct hadronic  $W/Z$  decays using variants of the boosted techniques mentioned previously. A combination of ATLAS analyses [76] searching in the dilepton, single-lepton-plus-jets, and fully hadronic final states for  $G^* \rightarrow VV$ , where  $V$  can represent either a  $W$  or  $Z$  boson, exclude gravitons with masses below 0.81 TeV, for a value of  $k/M_P = 1$ . CMS  $VV$  analyses with one boson decaying leptonically and the other hadronically [77] or both decaying hadronically [71] also provide cross section limits in the context of bulk gravitons; however, a maximum value of  $k/M_P = 0.5$  is presented, for which no mass

exclusion is possible using the full 8 TeV sample. Less restrictive limits in these models result from searching for  $G^* \rightarrow HH$  [78].

The lightest KK states are, in certain models, the partners of the top quark. For example, in 5D composite Higgs models these are colored states with charges  $Q = -1/3, 2/3$  and  $5/3$ , and masses expected to be below the TeV [64]. They can be either singly or pair-produced, and mainly decay into a combination of  $W/Z$  with top/bottom quarks [79]. Of particular note, the  $Q = 5/3$  state decays mainly into  $W^+t \rightarrow W^+W^+b$ , giving a pair of same-sign leptons in the final state. An analysis by ATLAS [81] searching in the lepton-plus-jets final state for evidence of pair production of the  $Q = 5/3$  state provides a lower mass limit of 840 GeV. Their analysis requiring in addition to a pair of same-sign leptons at least one  $b$ -tagged jet in the event [82] provides less stringent limits from pair production, and also from single production, the cross section for which is model-dependent [80]. A CMS same-sign dilepton analysis [83] searching for pair production of the  $Q = 5/3$  state excludes masses below 800 GeV. Both LHC experiments have searched for pair production of vector-like quarks  $T$  and  $B$  of charges  $Q = 2/3$  and  $-1/3$  respectively, assuming the allowable decays are  $T \rightarrow Wb/Zt/Ht$  and  $B \rightarrow Wt/Zb/Hb$ . In each case, it is assumed the branching fractions of the three decay modes sum to unity, but the individual branching fractions, which are model-dependent, are allowed to vary within this constraint. Depending on the values of the individual branching fractions, CMS obtains lower limits on the mass of the  $T$  [84] ( $B$  [85]) vector-like quark in the range of 720–920 GeV (740–900 GeV), while a summary [86] of the ATLAS searches provides lower limits on the  $T$  ( $B$ ) mass in the range of 715–950 GeV (575–813 GeV).

### III.2 Connection with Strongly-Coupled Models via the AdS/CFT Correspondence

The AdS/CFT correspondence [7] provides a connection between warped extra-dimensional models and strongly-coupled theories in ordinary 4D. Although the exact connection is only known for certain cases, the AdS/CFT techniques have been very useful to obtain, at the qualitative level, a 4D holographic description of the various phenomena in warped extra-dimensional models [8].

The connection goes as follows. The physics of the bulk AdS<sub>5</sub> models can be interpreted as that of a 4D conformal field theory (CFT) which is strongly coupled. The extra-dimensional coordinate  $y$  plays the role of the renormalization scale  $\mu$  of the CFT by means of the identification  $\mu \equiv ke^{-ky}$ . Therefore the UV-boundary corresponds in the CFT to a UV cut-off scale at  $\Lambda_{UV} = k \sim M_P$ , breaking explicitly conformal invariance, while the IR-boundary can be interpreted as a spontaneous breaking of the conformal symmetry at energies  $ke^{-k\pi R} \sim \text{TeV}$ . Fields localized on the UV-boundary are elementary fields external to the CFT, while fields localized on the IR-boundary and KK states corresponds to composite resonances of the CFT. Furthermore, local gauge symmetries in the 5D models,  $G_5$ ,

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### Extra Dimensions

correspond to global symmetries of the CFT, while the UV-boundary symmetry can be interpreted as a gauging of the subgroup  $\mathcal{H}_{UV}$  of  $\mathcal{G}_5$  in the CFT. Breaking gauge symmetries by IR-boundary conditions corresponds to the spontaneous breaking  $\mathcal{G}_5 \rightarrow \mathcal{H}_{IR}$  in the CFT at energies  $\sim ke^{-k\pi R}$ . Using this correspondence one can easily derive the 4D massless spectrum of the compactified AdS<sub>5</sub> models. One also has the identification  $k^3/M_5^3 \approx 16\pi^2/N^2$  and  $g_5^2 k \approx 16\pi^2/N^r$  ( $r = 1$  or  $2$  for CFT fields in the fundamental or adjoint representation of the gauge group), where  $N$  plays the role of the number of colors of the CFT. Therefore the weak-coupling limit in AdS<sub>5</sub> corresponds to a large- $N$  expansion in the CFT.

Following the above AdS/CFT dictionary one can understand the RS solution to the hierarchy problem from a 4D viewpoint. The equivalent 4D model is a CFT with a TeV mass-gap and a Higgs emerging as a composite state. In the particular case where the Higgs is the fifth-component of the gauge-boson,  $A_5$  [87], this corresponds to models, similar to those proposed in Ref. [88], where the Higgs is a composite pseudo-Goldstone boson arising from the spontaneous breaking  $\mathcal{G}_5 \rightarrow \mathcal{H}_{IR}$  in the CFT. The AdS/CFT dictionary tells us that KK states must behave as composite resonances. For example, if the SM gauge bosons propagate in the 5D bulk, the lowest KK  $SU(2)_L$ -gauge boson must have properties similar to those of the Techni-rho  $\rho_T$  [89] with a coupling to longitudinal  $W/Z$  bosons given by  $g_5\sqrt{k} \approx g_{\rho T}$ , while the coupling to elementary fermions is  $g^2/\sqrt{g_5^2 k} \approx g^2 F_{\rho T}/M_{\rho T}$ .

Fermions in compactified AdS<sub>5</sub> also have a simple 4D holographic interpretation. The 4D massless mode described in Sec. III.1 corresponds to an external fermion  $\psi_i$  linearly coupled to a fermionic CFT operator  $\mathcal{O}_i$ :  $\mathcal{L}_{\text{int}} = \lambda_i \bar{\psi}_i \mathcal{O}_i + h.c.$ . The dimension of the operator  $\mathcal{O}_i$  is related to the 5D fermion mass according to  $\text{Dim}[\mathcal{O}_i] = |c_f + 1/2| - 1$ . Therefore, by varying  $c_f$  one varies  $\text{Dim}[\mathcal{O}_i]$ , making the coupling  $\lambda_i$  irrelevant ( $c_f > 1/2$ ), marginal ( $c_f = 1/2$ ) or relevant ( $c_f < 1/2$ ). When irrelevant, the coupling is exponentially suppressed at low energies, and then the coupling of  $\psi_i$  to the CFT (and eventually to the composite Higgs) is very small. When relevant, the coupling grows in the IR and become as large as  $g_5$  (in units of  $k$ ), meaning that the fermion is as strongly coupled as the CFT states [62]. In this latter case  $\psi_i$  behaves as a composite fermion.

### III.3 Flat Extra Dimensions

Models with quantum-gravity at the TeV scale, as in the ADD scenario, can have extra (flat) dimensions of  $1/\text{TeV}$  size, as happens in string scenarios [90]. All SM fields may propagate in these extra dimensions, leading to the possibility of observing their corresponding KK states.

A simple example is to assume that the SM gauge bosons propagate in a flat five-dimensional orbifold  $S^1/Z_2$  of radius  $R$ , with the fermions localized on a 4D boundary. The KK gauge bosons behave as sequential SM gauge bosons with a coupling to fermions enhanced by a factor  $\sqrt{2}$  [90]. The

experimental limits on such sequential gauge bosons could therefore be recast as limits on KK gauge bosons. Such an interpretation of the ATLAS 7 TeV dilepton analysis [91] yielded the bound  $1/R > 4.16$  TeV, while a CMS 8 TeV search with a lepton and missing transverse energy in the final state [92] give  $1/R > 3.4$  TeV. Indirect bounds from LEP2 require however  $1/R \gtrsim 6$  TeV [93,65].

An alternative scenario, known as Universal Extra Dimensions (UED) [94], assumes that all SM fields propagate universally in a flat orbifold  $S^1/Z_2$  with an extra  $Z_2$  parity, called KK-parity, that interchanges the two boundaries. In this case, the lowest KK state is stable and is a Dark Matter candidate. At colliders, the KK particles would have to be created in pairs, and would then cascade decay to the lightest KK particle (LKP), which would be stable and escape detection. Experimental signatures, such as jets or leptons and  $\cancel{E}_T$ , would be similar to those of typical  $R$ -parity conserving SUSY searches. Theoretical studies of the trilepton final state [95] suggest a potential bound from the LHC at 8 TeV with  $20 \text{ fb}^{-1}$  of  $1/R \gtrsim 1.3$  TeV for  $\Lambda R = 10$ , where  $\Lambda$  is the cut-off scale of the model. The experimental searches have not yet been interpreted in the general UED scenario; for example, the ATLAS trilepton analysis [96] of their full 8 TeV dataset provides upper limits on the visible cross section for new physics that could be utilized to determine UED limits.

Experimental limits have been provided on two specific UED models which include KK parity violation. In one case, KK parity is violated by gravitational interactions [97], and the LKP can decay via  $\gamma^* \rightarrow \gamma + G$ . Beginning with strong production of a pair of KK quarks and/or gluons [98,99], the final state would be  $\gamma\gamma + \cancel{E}_T + X$ . Using their full 7 TeV datasets, ATLAS [100] and CMS [101] each determine a limit of  $1/R \gtrsim 1.4$  TeV for  $\Lambda R = 20$ . In a second model, that involves two UEDs, the breaking of the KK parity allows the decay of the KK photon to  $t\bar{t}$  [102]. The ATLAS vector-like quark search [86] (same-sign dilepton plus b-jet analysis [82]), applied to search for a  $t\bar{t}t\bar{t}$  final state for evidence of pair-produced KK photons, excludes KK masses below 1.12 TeV (0.96 TeV) in this model.

Finally, realistic models of electroweak symmetry breaking can also be constructed with flat extra spatial dimensions, similarly to those in the warped case, requiring, however, the presence of sizeable boundary kinetic terms [103]. There is also the possibility of breaking supersymmetry by boundary conditions [104]. Models of this type could explain naturally the presence of a Higgs boson lighter than  $M_D \sim \text{TeV}$  [105].

### References

1. For a comprehensive collection of the original papers see, “Modern Kaluza-Klein Theories”, edited by T. Appelquist *et al.*, Addison-Wesley (1987).
2. N. Arkani-Hamed *et al.*, Phys. Lett. **B429**, 263 (1998).
3. L. Randall and R. Sundrum, Phys. Rev. Lett. **83**, 3370 (1999).

4. H. Davoudiasl *et al.*, Phys. Lett. **B473**, 43 (2000); A. Pomarol, Phys. Lett. **B486**, 153 (2000).
5. S. Chang *et al.*, Phys. Rev. **D62**, 084025 (2000).
6. T. Gherghetta and A. Pomarol, Nucl. Phys. **B586**, 141 (2000).
7. J.M. Maldacena, Adv. Theor. Math. Phys. **2**, 231 (1998); E. Witten, Adv. Theor. Math. Phys. **2**, 253 (1998); S.S. Gubser *et al.*, Phys. Lett. **B428**, 105 (1998).
8. N. Arkani-Hamed *et al.*, JHEP **0108**, 017 (2001).
9. ATLAS public results are available on WWW at <https://twiki.cern.ch/twiki/bin/view/AtlasPublic>.
10. CMS public results are available on WWW at [cms-results.web.cern.ch/cms-results/public-results/publications](https://cms-results.web.cern.ch/cms-results/public-results/publications).
11. Z. Chacko *et al.*, JHEP **0007**, 036 (2000).
12. N. Arkani-Hamed *et al.*, Phys. Rev. **D59**, 086004 (1999).
13. For a review see for example, R. Rattazzi, hep-ph/0607055 (2006); I. Antoniadis, Yellow report CERN-2002-002 (2002).
14. I. Antoniadis *et al.*, Phys. Lett. **B436**, 257 (1998).
15. G.F. Giudice *et al.*, Nucl. Phys. **B544**, 3 (1999).
16. For the case of two extra dimensions, see for example, N. Arkani-Hamed *et al.*, Phys. Rev. **D62**, 105002 (2000).
17. E.G. Adelberger *et al.*, Prog. in Part. Nucl. Phys. **62**, 102 (2009); J. Murata and S. Tanaka, Class. Quantum Grav. **32**, 033001 (2015).
18. C. Hanhart *et al.*, Phys. Lett. **B509**, 1 (2001).
19. S. Hannestad and G.G. Raffelt, Phys. Rev. **D67**, 125008 (2003).
20. L.J. Hall and D. Tucker-Smith, Phys. Rev. **D60**, 085008 (1999).
21. E.A. Mirabelli *et al.*, Phys. Rev. Lett. **82**, 2236 (1999).
22. T. Han *et al.*, Phys. Rev. **D59**, 105006 (1999).
23. CMS Collab., Eur. Phys. J. **C75**, 235 (2015).
24. ATLAS Collab., Eur. Phys. J. **C75**, 299 (2015).
25. ATLAS Collab., Phys. Rev. **D91**, 012008 (2015).
26. CMS Collab., \tt arXiv:1410.8812, submitted to PLB.
27. See for example S. Cullen *et al.*, Phys. Rev. **D62**, 055012 (2000).
28. CMS Collab., CMS-PAS-EXO-15-001.
29. CMS Collab., Phys. Rev. **D91**, 052009 (2015).
30. ATLAS Collab., Phys. Rev. **D91**, 052007 (2015).
31. ATLAS Collab., ATLAS-CONF-2015-042.
32. J.L. Hewett, Phys. Rev. Lett. **82**, 4765 (1999).
33. G.F. Giudice and A. Strumia, Nucl. Phys. **B663**, 377 (2003).
34. CMS Collab., Phys. Lett. **B746**, 79 (2015).
35. ATLAS Collab., Phys. Rev. Lett. **114**, 221802 (2015).
36. ATLAS Collab., Eur. Phys. J. **C74**, 3134 (2014).
37. CMS Collab., JHEP **1504**, 025 (2015).
38. S.B. Giddings and S. Thomas, Phys. Rev. **D65**, 056010 (2002); S. Dimopoulos and G. Landsberg, Phys. Rev. Lett. **87**, 161602 (2001); for a review see for example, P. Kanti, Int. J. Mod. Phys. **A19**, 4899 (2004).
39. R.C. Myers and M.J. Perry, Ann. Phys. **172**, 304 (1986).
40. ATLAS Collab., ATLAS-CONF-2015-046.
41. ATLAS Collab., JHEP **1408**, 103 (2014).
42. ATLAS Collab., ATLAS-CONF-2015-043.
43. ATLAS Collab., JHEP **1507**, 032 (2015).
44. ATLAS Collab., Phys. Rev. **D88**, 072001 (2013).
45. CMS Collab., JHEP **1307**, 178 (2013).
46. P. Meade and L. Randall, JHEP **0805**, 003 (2008).
47. ATLAS Collab., Phys. Lett. **B728**, 562 (2013).
48. ATLAS Collab., Phys. Rev. Lett. **112**, 091804 (2013).
49. S. Dimopoulos and R. Emparan, Phys. Lett. **B526**, 393 (2002).
50. W.D. Goldberger and M.B. Wise, Phys. Rev. Lett. **83**, 4922 (1999); J. Garriga, A. Pomarol, Phys. Lett. **B560**, 91 (2003).
51. See talk by R. Rattazzi at Planck 2010, CERN; B. Bellazzini *et al.*, Eur. Phys. J. **C74**, 2790 (2014); F. Coradeschi *et al.*, JHEP **1311**, 057 (2013); E. Megias and O. Pujolas, JHEP **1408**, 081 (2014).
52. For a review see for example, H. Davoudiasl *et al.*, New J. Phys. **12**, 075011 (2010); T. Gherghetta, arXiv:1008.2570.
53. Y. Grossman and M. Neubert, Phys. Lett. **B474**, 361 (2000).
54. S.J. Huber and Q. Shafi, Phys. Lett. **B498**, 256 (2001).
55. A. Delgado *et al.*, JHEP **0001**, 030 (2000).
56. K. Agashe *et al.*, Phys. Rev. **D71**, 016002 (2005); for a recent analysis see for example, M. Bauer *et al.*, JHEP **1009**, 017 (2010).
57. For a review see for example, A. Pomarol, Int. J. Mod. Phys. **A24**, 61 (2009).
58. See, for example, PDG review of Higgs boson in this Review.
59. C. Csaki *et al.*, Phys. Rev. Lett. **92**, 101802 (2004); for a review see for example, C. Csaki *et al.*, hep-ph/0510275.
60. Y. Hosotani, Phys. Lett. **B126**, 309 (1983).
61. K. Agashe *et al.*, JHEP **0308**, 050 (2003).
62. K. Agashe *et al.*, Nucl. Phys. **B719**, 165 (2005); for a review see for example, R. Contino, arXiv:1005.4269.
63. G.F. Giudice *et al.*, JHEP **0706**, 045 (2007).
64. R. Contino *et al.*, Phys. Rev. **D75**, 055014 (2007).
65. R. Barbieri *et al.*, Nucl. Phys. **B703**, 127 (2004).
66. CMS Collab., JHEP **1504**, 025 (2015).
67. ATLAS Collab., Phys. Rev. **D90**, 052005 (2014).
68. ATLAS Collab., Phys. Rev. **D92**, 032004 (2015).
69. CMS Collab., CMS-PAS-EXO-12-045.
70. CMS Collab., arXiv:1503.04114, submitted to PLB.
71. CMS Collab., JHEP **1408**, 173 (2014).
72. K. Agashe *et al.*, Phys. Rev. **D76**, 036006 (2007).
73. CMS Collab., arXiv:1506.03062, submitted to PRD.
74. ATLAS Collab., JHEP **1508**, 148 (2015).
75. See, for example, PDG review of  $W'$  boson searches in this Review.
76. ATLAS Collab., ATLAS-CONF-2015-045.
77. CMS Collab., JHEP **1408**, 174 (2014).
78. ATLAS Collab., arXiv:1506.00285, accepted by EPJC.
79. R. Contino and G. Servant, JHEP **0806**, 026 (2008); J.A. Aguilar-Saavedra, JHEP **0911**, 030 (2009); J. Mrazek and A. Wulzer, Phys. Rev. **D81**, 075006 (2010); G. Dissertori *et al.*, JHEP **1009**, 019 (2010).

# Searches Particle Listings

## Extra Dimensions

80. A. De Simone *et al.*, JHEP **1304**, 004 (2013).
81. ATLAS Collab., Phys. Rev. **D91**, 112011 (2015).
82. ATLAS Collab., arXiv:1504.04605, submitted to JHEP.
83. CMS Collab., Phys. Rev. Lett. **112**, 171801 (2014).
84. CMS Collab., arXiv:1509.04177, submitted to PRD.
85. CMS Collab., arXiv:1507.07129, submitted to PRD.
86. ATLAS Collab., JHEP **1508**, 105 (2015).
87. R. Contino, Y. Nomura and A. Pomarol, Nucl. Phys. **B671**, 148 (2003).
88. H. Georgi *et al.*, Phys. Lett. **B143**, 152 (1984); D.B. Kaplan *et al.*, Phys. Lett. **B136**, 183 (1984).
89. See, for example, PDG review of ‘*Dynamical Electroweak Symmetry Breaking*’ in this Review.
90. See for example, I. Antoniadis and K. Benakli, Int. J. Mod. Phys. **A15**, 4237 (2000).
91. ATLAS Collab., JHEP **1211**, 138 (2012).
92. CMS Collab., Phys. Rev. **D91**, 092005 (2015).
93. K. Cheung and G. L. Landsberg, Phys. Rev. **D65**, 076003 (2002).
94. T. Appelquist *et al.*, Phys. Rev. **D64**, 035002 (2001); for a review see for example, A. Datta *et al.*, New J. Phys. **12**, 075017 (2010).
95. A. Belyaev *et al.*, JHEP **1306**, 080 (2013).
96. ATLAS Collab., JHEP **1508**, 138 (2015).
97. C. Macesanu, *et al.*, Phys. Lett. **B546**, 253 (2002).
98. C. Macesanu *et al.*, Phys. Rev. **D66**, 015009 (2002).
99. H.C. Cheng *et al.*, Phys. Rev. **D66**, 036005 (2002).
100. ATLAS Collab., Phys. Lett. **B718**, 411 (2012).
101. CMS Collab., JHEP **1303**, 111 (2013).
102. G. Burdman *et al.*, Phys. Rev. **D74**, 075008 (2006).
103. For a review see for example, G. Panico *et al.*, JHEP **1102**, 103 (2011).
104. J. Scherk and J.H. Schwarz, Phys. Lett. **B82**, 60 (1979).
105. See for example, A. Pomarol and M. Quiros, Phys. Lett. **B438**, 255 (1998); I. Antoniadis *et al.*, Nucl. Phys. **B544**, 503 (1999); R. Barbieri *et al.*, Phys. Rev. **D63**, 105007 (2001).

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 Limits on Kaluza-Klein Gluons in Warped Extra Dimensions

### Limits on $R$ from Deviations in Gravitational Force Law

This section includes limits on the size of extra dimensions from deviations in the Newtonian ( $1/r^2$ ) gravitational force law at short distances. Deviations are parametrized by a gravitational potential of the form  $V = -(G m m'/r) [1 + \alpha \exp(-r/R)]$ . For  $\delta$  toroidal extra dimensions of equal size,  $\alpha = 8\delta/3$ . Quoted bounds are for  $\delta = 2$  unless otherwise noted.

VALUE ( $\mu\text{m}$ )	CL%	DOCUMENT ID	COMMENT
< 30	95	1 KAPNER 07	Torsion pendulum
• • • We do not use the following data for averages, fits, limits, etc. • • •			
		2 XU 13	Nuclei properties
		3 BEZERRA 11	Torsion oscillator
		4 SUSHKOV 11	Torsion pendulum
		5 BEZERRA 10	Microcantilever
		6 MASUDA 09	Torsion pendulum
		7 GERACI 08	Microcantilever
		8 TRENKEL 08	Newton's constant
		9 DECCA 07A	Torsion oscillator

< 47	95	10 TU 07	Torsion pendulum
		11 SMULLIN 05	Microcantilever
<130	95	12 HOYLE 04	Torsion pendulum
		13 CHIAVERINI 03	Microcantilever
$\lesssim 200$	95	14 LONG 03	Microcantilever
<190	95	15 HOYLE 01	Torsion pendulum
		16 HOSKINS 85	Torsion pendulum

- 1 KAPNER 07 search for new forces, probing a range of  $\alpha \simeq 10^{-3}$ – $10^5$  and length scales  $R \simeq 10$ – $1000 \mu\text{m}$ . For  $\delta = 1$  the bound on  $R$  is  $44 \mu\text{m}$ . For  $\delta = 2$ , the bound is expressed in terms of  $M_*$ , here translated to a bound on the radius. See their Fig. 6 for details on the bound.
- 2 XU 13 obtain constraints on non-Newtonian forces with strengths  $|\alpha| \simeq 10^{34}$ – $10^{36}$  and length scales  $R \simeq 1$ – $10 \text{ fm}$ . See their Fig. 4 for more details. These constraints do not place limits on the size of extra flat dimensions.
- 3 BEZERRA 11 obtain constraints on non-Newtonian forces with strengths  $10^{11} \lesssim |\alpha| \lesssim 10^{18}$  and length scales  $R = 30$ – $1260 \text{ nm}$ . See their Fig. 2 for more details. These constraints do not place limits on the size of extra flat dimensions.
- 4 SUSHKOV 11 obtain improved limits on non-Newtonian forces with strengths  $10^7 \lesssim |\alpha| \lesssim 10^{11}$  and length scales  $0.4 \mu\text{m} < R < 4 \mu\text{m}$  (95% CL). See their Fig. 2. These bounds do not place limits on the size of extra flat dimensions. However, a model dependent bound of  $M_* > 70 \text{ TeV}$  is obtained assuming gauge bosons that couple to baryon number also propagate in  $(4 + \delta)$  dimensions.
- 5 BEZERRA 10 obtain improved constraints on non-Newtonian forces with strengths  $10^{19} \lesssim |\alpha| \lesssim 10^{29}$  and length scales  $R = 1.6$ – $14 \text{ nm}$  (95% CL). See their Fig. 1. This bound does not place limits on the size of extra flat dimensions.
- 6 MASUDA 09 obtain improved constraints on non-Newtonian forces with strengths  $10^9 \lesssim |\alpha| \lesssim 10^{11}$  and length scales  $R = 1.0$ – $2.9 \mu\text{m}$  (95% CL). See their Fig. 3. This bound does not place limits on the size of extra flat dimensions.
- 7 GERACI 08 obtain improved constraints on non-Newtonian forces with strengths  $|\alpha| > 14,000$  and length scales  $R = 5$ – $15 \mu\text{m}$ . See their Fig. 9. This bound does not place limits on the size of extra flat dimensions.
- 8 TRENKEL 08 uses two independent measurements of Newton's constant  $G$  to constrain new forces with strength  $|\alpha| \simeq 10^{-4}$  and length scales  $R = 0.02$ – $1 \text{ m}$ . See their Fig. 1. This bound does not place limits on the size of extra flat dimensions.
- 9 DECCA 07A search for new forces and obtain bounds in the region with strengths  $|\alpha| \simeq 10^{13}$ – $10^{18}$  and length scales  $R = 20$ – $86 \text{ nm}$ . See their Fig. 6. This bound does not place limits on the size of extra flat dimensions.
- 10 TU 07 search for new forces probing a range of  $|\alpha| \simeq 10^{-1}$ – $10^5$  and length scales  $R \simeq 20$ – $1000 \mu\text{m}$ . For  $\delta = 1$  the bound on  $R$  is  $53 \mu\text{m}$ . See their Fig. 3 for details on the bound.
- 11 SMULLIN 05 search for new forces, and obtain bounds in the region with strengths  $\alpha \simeq 10^3$ – $10^8$  and length scales  $R = 6$ – $20 \mu\text{m}$ . See their Figs. 1 and 16 for details on the bound. This work does not place limits on the size of extra flat dimensions.
- 12 HOYLE 04 search for new forces, probing  $\alpha$  down to  $10^{-2}$  and distances down to  $10 \mu\text{m}$ . Quoted bound on  $R$  is for  $\delta = 2$ . For  $\delta = 1$ , bound goes to  $160 \mu\text{m}$ . See their Fig. 34 for details on the bound.
- 13 CHIAVERINI 03 search for new forces, probing  $\alpha$  above  $10^4$  and  $\lambda$  down to  $3 \mu\text{m}$ , finding no signal. See their Fig. 4 for details on the bound. This bound does not place limits on the size of extra flat dimensions.
- 14 LONG 03 search for new forces, probing  $\alpha$  down to 3, and distances down to about  $10 \mu\text{m}$ . See their Fig. 4 for details on the bound.
- 15 HOYLE 01 search for new forces, probing  $\alpha$  down to  $10^{-2}$  and distances down to  $20 \mu\text{m}$ . See their Fig. 4 for details on the bound. The quoted bound is for  $\alpha \geq 3$ .
- 16 HOSKINS 85 search for new forces, probing distances down to  $4 \text{ mm}$ . See their Fig. 13 for details on the bound. This bound does not place limits on the size of extra flat dimensions.

### Limits on $R$ from On-Shell Production of Gravitons: $\delta = 2$

This section includes limits on on-shell production of gravitons in collider and astrophysical processes. Bounds quoted are on  $R$ , the assumed common radius of the flat extra dimensions, for  $\delta = 2$  extra dimensions. Studies often quote bounds in terms of derived parameter; experiments are actually sensitive to the masses of the KK gravitons:  $m_{\vec{n}} = |\vec{n}|/R$ . See the Review on “Extra Dimensions” for details. Bounds are given in  $\mu\text{m}$  for  $\delta = 2$ .

VALUE ( $\mu\text{m}$ )	CL%	DOCUMENT ID	TECN	COMMENT
< 15	95	1 KHACHATRYAN 15AL	CMS	$p\bar{p} \rightarrow jG$
< 0.00016	95	2 HANNESTAD 03		Neutron star heating
• • • We do not use the following data for averages, fits, limits, etc. • • •				
	95	3 AAD 15CS	ATLS	$p\bar{p} \rightarrow \gamma G$
< 25	95	4 AAD 13AD	ATLS	$p\bar{p} \rightarrow jG$
< 127	95	5 AAD 13C	ATLS	$p\bar{p} \rightarrow \gamma G$
< 34.4	95	6 AAD 13D	ATLS	$p\bar{p} \rightarrow jj$
< 0.0087	95	7 AJELLO 12	FLAT	Neutron star $\gamma$ sources
< 23	95	8 CHATRCHYAN 12AP	CMS	$p\bar{p} \rightarrow jG$
< 92	95	9 AAD 11S	ATLS	$p\bar{p} \rightarrow jG$
< 72	95	10 CHATRCHYAN 11U	CMS	$p\bar{p} \rightarrow jG$
< 245	95	11 AALTONEN 08AC	CDF	$p\bar{p} \rightarrow \gamma G, jG$
< 615	95	12 ABAZOV 08S	D0	$p\bar{p} \rightarrow \gamma G$
< 0.916	95	13 DAS 08		Supernova cooling
< 350	95	14 ABULENCIA, A 06	CDF	$p\bar{p} \rightarrow jG$
< 270	95	15 ABDALLAH 05B	DLPH	$e^+e^- \rightarrow \gamma G$
< 210	95	16 ACHARD 04E	L3	$e^+e^- \rightarrow \gamma G$
< 480	95	17 ACOSTA 04C	CDF	$p\bar{p} \rightarrow jG$
< 0.00038	95	18 CASSE 04		Neutron star $\gamma$ sources
< 610	95	19 ABAZOV 03	D0	$p\bar{p} \rightarrow jG$
< 0.96	95	20 HANNESTAD 03		Supernova cooling

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< 0.096	95	21	HANNESTAD	03	Diffuse $\gamma$ background
< 0.051	95	22	HANNESTAD	03	Neutron star $\gamma$ sources
< 300	95	23	HEISTER	03c ALEP	$e^+e^- \rightarrow \gamma G$
		24	FAIRBAIRN	01	Cosmology
< 0.66	95	25	HANHART	01	Supernova cooling
		26	CASSISI	00	Red giants
<1300	95	27	ACCIARRI	99s L3	$e^+e^- \rightarrow ZG$

- 1 KHACHATRYAN 15AL search for  $pp \rightarrow jG$ , using  $19.7 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 8 \text{ TeV}$  to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on  $R$  is derived. See their Table 7 for bounds on all  $\delta \leq 6$ .
- 2 HANNESTAD 03 obtain a limit on  $R$  from the heating of old neutron stars by the surrounding cloud of trapped KK gravitons. Limits for all  $\delta \leq 7$  are given in their Tables V and VI. These limits supersede those in HANNESTAD 02.
- 3 AAD 15Cs search for  $pp \rightarrow \gamma G$ , using  $20.3 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 8 \text{ TeV}$  to place lower limits on  $M_D$  for two to six extra dimensions (see their Fig. 18).
- 4 AAD 13AD search for  $pp \rightarrow jG$ , using  $4.7 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$  to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on  $R$  is derived. See their Table 8 for bounds on all  $\delta \leq 6$ .
- 5 AAD 13c search for  $pp \rightarrow \gamma G$ , using  $4.6 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$  to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on  $R$  is derived.
- 6 AAD 13b search for the dijet decay of quantum black holes in  $4.8 \text{ fb}^{-1}$  of data produced in  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  to place bounds on  $M_D$  for two to seven extra dimensions, from which these bounds on  $R$  are derived. Limits on  $M_D$  for all  $\delta \leq 7$  are given in their Table 3.
- 7 AJELLO 12 obtain a limit on  $R$  from the gamma-ray emission of point  $\gamma$  sources that arise from the photon decay of KK gravitons which are gravitationally bound around neutron stars. Limits for all  $\delta \leq 7$  are given in their Table 7.
- 8 CHATRCHYAN 12AP search for  $pp \rightarrow jG$ , using  $5.0 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$  to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on  $R$  is derived. See their Table 7 for bounds on all  $\delta \leq 6$ .
- 9 AAD 11s search for  $pp \rightarrow jG$ , using  $33 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$ , to place bounds on  $M_D$  for two to four extra dimensions, from which these bounds on  $R$  are derived. See their Table 3 for bounds on all  $\delta \leq 4$ .
- 10 CHATRCHYAN 11u search for  $pp \rightarrow jG$ , using  $36 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 7 \text{ TeV}$ , to place bounds on  $M_D$  for two to six extra dimensions, from which these bounds on  $R$  are derived. See their Table 3 for bounds on all  $\delta \leq 6$ .
- 11 AALTONEN 08AC search for  $p\bar{p} \rightarrow \gamma G$  and  $p\bar{p} \rightarrow jG$  at  $\sqrt{s} = 1.96 \text{ TeV}$  with  $2.0 \text{ fb}^{-1}$  and  $1.1 \text{ fb}^{-1}$  respectively, in order to place bounds on the fundamental scale and size of the extra dimensions. See their Table III for limits on all  $\delta \leq 6$ .
- 12 ABAZOV 08s search for  $p\bar{p} \rightarrow \gamma G$ , using  $1 \text{ fb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$  to place bounds on  $M_D$  for two to eight extra dimensions, from which these bounds on  $R$  are derived. See their paper for intermediate values of  $\delta$ .
- 13 DAS 08 obtain a limit on  $R$  from Kaluza-Klein graviton cooling of SN1987A due to plasmon-plasmon annihilation.
- 14 ABULENCIA 06 search for  $p\bar{p} \rightarrow jG$  using  $368 \text{ pb}^{-1}$  of data at  $\sqrt{s} = 1.96 \text{ TeV}$ . See their Table II for bounds for all  $\delta \leq 6$ .
- 15 ABDALLAH 05B search for  $e^+e^- \rightarrow \gamma G$  at  $\sqrt{s} = 180\text{--}209 \text{ GeV}$  to place bounds on the size of extra dimensions and the fundamental scale. Limits for all  $\delta \leq 6$  are given in their Table 6. These limits supersede those in ABREU 002.
- 16 ACHARD 04E search for  $e^+e^- \rightarrow \gamma G$  at  $\sqrt{s} = 189\text{--}209 \text{ GeV}$  to place bounds on the size of extra dimensions and the fundamental scale. See their Table 8 for limits with  $\delta \leq 8$ . These limits supersede those in ACCIARRI 99R.
- 17 ACOSTA 04C search for  $p\bar{p} \rightarrow jG$  at  $\sqrt{s} = 1.8 \text{ TeV}$  to place bounds on the size of extra dimensions and the fundamental scale. See their paper for bounds on  $\delta = 4, 6$ .
- 18 CASSE 04 obtain a limit on  $R$  from the gamma-ray emission of point  $\gamma$  sources that arises from the photon decay of gravitons around newly born neutron stars, applying the technique of HANNESTAD 03 to neutron stars in the galactic bulge. Limits for all  $\delta \leq 7$  are given in their Table I.
- 19 ABAZOV 03 search for  $p\bar{p} \rightarrow jG$  at  $\sqrt{s} = 1.8 \text{ TeV}$  to place bounds on  $M_D$  for 2 to 7 extra dimensions, from which these bounds on  $R$  are derived. See their paper for bounds on intermediate values of  $\delta$ . We quote results without the approximate NLO scaling introduced in the paper.
- 20 HANNESTAD 03 obtain a limit on  $R$  from graviton cooling of supernova SN1987A. Limits for all  $\delta \leq 7$  are given in their Tables V and VI.
- 21 HANNESTAD 03 obtain a limit on  $R$  from gravitons emitted in supernovae and which subsequently decay, contaminating the diffuse cosmic  $\gamma$  background. Limits for all  $\delta \leq 7$  are given in their Tables V and VI. These limits supersede those in HANNESTAD 02.
- 22 HANNESTAD 03 obtain a limit on  $R$  from gravitons emitted in two recent supernovae and which subsequently decay, creating point  $\gamma$  sources. Limits for all  $\delta \leq 7$  are given in their Tables V and VI. These limits are corrected in the published erratum.
- 23 HEISTER 03c use the process  $e^+e^- \rightarrow \gamma G$  at  $\sqrt{s} = 189\text{--}209 \text{ GeV}$  to place bounds on the size of extra dimensions and the scale of gravity. See their Table 4 for limits with  $\delta \leq 6$  for derived limits on  $M_D$ .
- 24 FAIRBAIRN 01 obtains bounds on  $R$  from over production of KK gravitons in the early universe. Bounds are quoted in paper in terms of fundamental scale of gravity. Bounds depend strongly on temperature of QCD phase transition and range from  $R < 0.13 \mu\text{m}$  to  $0.001 \mu\text{m}$  for  $\delta=2$ ; bounds for  $\delta=3,4$  can be derived from Table I in the paper.
- 25 HANHART 01 obtain bounds on  $R$  from limits on graviton cooling of supernova SN1987a using numerical simulations of proto-neutron star neutrino emission.
- 26 CASSISI 00 obtain rough bounds on  $M_D$  (and thus  $R$ ) from red giant cooling for  $\delta=2,3$ . See their paper for details.
- 27 ACCIARRI 99s search for  $e^+e^- \rightarrow ZG$  at  $\sqrt{s}=189 \text{ GeV}$ . Limits on the gravity scale are found in their Table 2, for  $\delta \leq 4$ .

### Mass Limits on $M_{TT}$

This section includes limits on the cut-off mass scale,  $M_{TT}$ , of dimension-8 operators from KK graviton exchange in models of large extra dimensions. Ambiguities in the UV-divergent summation are absorbed into the parameter  $\lambda$ , which is taken to be  $\lambda = \pm 1$  in the following analyses. Bounds for  $\lambda = -1$  are shown in parenthesis after the bound for  $\lambda = +1$ , if appropriate. Different papers use slightly different definitions of the mass scale. The definition used here is related to another popular convention by  $M_{TT}^4 = (2/\pi) \Lambda_4^4$ , as discussed in the above Review on "Extra Dimensions."

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 6.3	95	1 KHACHATRYAN..15J	CMS	$pp \rightarrow$ dijet, ang. distrib.
>20.6	(> 15.7)	2 GIUDICE 03	RVUE	Dim-6 operators

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 3.7	95	3 KHACHATRYAN..15AE	CMS	$pp \rightarrow e^+e^-, \mu^+\mu^-$
> 3.8	95	4 AAD 14BE	ATLS	$pp \rightarrow e^+e^-, \mu^+\mu^-$
> 2.94	(>2.52)	5 AAD 13AS	ATLS	$pp \rightarrow \gamma\gamma$
> 3.2	95	6 AAD 13E	ATLS	$pp \rightarrow e^+e^-, \mu^+\mu^-, \gamma\gamma$
> 2.66	(>2.27)	7 AAD 12Y	ATLS	$pp \rightarrow \gamma\gamma$
		8 BAAK 12	RVUE	Electroweak
> 2.86	95	9 CHATRCHYAN12J	CMS	$pp \rightarrow e^+e^-, \mu^+\mu^-$
> 2.84	(>2.41)	10 CHATRCHYAN12R	CMS	$pp \rightarrow \gamma\gamma$
> 0.90	(>0.92)	11 AARON 11c	H1	$e^\pm p \rightarrow e^\pm X$
> 1.74	(>1.71)	12 CHATRCHYAN11A	CMS	$pp \rightarrow \gamma\gamma$
> 1.48	95	13 ABAZOV 09AE	D0	$p\bar{p} \rightarrow$ dijet, ang. distrib.
> 1.45	95	14 ABAZOV 09D	D0	$p\bar{p} \rightarrow e^+e^-, \gamma\gamma$
> 1.1	(> 1.0)	15 SCHAEEL 07A	ALEP	$e^+e^- \rightarrow e^+e^-$
> 0.898	(> 0.998)	16 ABDALLAH 06C	DLPH	$e^+e^- \rightarrow \ell^+\ell^-$
> 0.853	(> 0.939)	17 GERDES 06		$p\bar{p} \rightarrow e^+e^-, \gamma\gamma$
> 0.96	(> 0.93)	18 ABAZOV 05v	D0	$p\bar{p} \rightarrow \mu^+\mu^-$
> 0.78	(> 0.79)	19 CHEKANOV 04B	ZEUS	$e^\pm p \rightarrow e^\pm X$
> 0.805	(> 0.956)	20 ABBIENDI 03D	OPAL	$e^+e^- \rightarrow \gamma\gamma$
> 0.7	(> 0.7)	21 ACHARD 03D	L3	$e^+e^- \rightarrow ZZ$
> 0.82	(> 0.78)	22 ADLOFF 03	H1	$e^\pm p \rightarrow e^\pm X$
> 1.28	(> 1.25)	23 GIUDICE 03	RVUE	
> 0.80	(> 0.85)	24 HEISTER 03C	ALEP	$e^+e^- \rightarrow \gamma\gamma$
> 0.84	(> 0.99)	25 ACHARD 02D	L3	$e^+e^- \rightarrow \gamma\gamma$
> 1.2	(> 1.1)	26 ABBOTT 01	D0	$p\bar{p} \rightarrow e^+e^-, \gamma\gamma$
> 0.60	(> 0.63)	27 ABBIENDI 00R	OPAL	$e^+e^- \rightarrow \mu^+\mu^-$
> 0.63	(> 0.50)	27 ABBIENDI 00R	OPAL	$e^+e^- \rightarrow \tau^+\tau^-$
> 0.68	(> 0.61)	27 ABBIENDI 00R	OPAL	$e^+e^- \rightarrow \mu^+\mu^-, \tau^+\tau^-$
		28 ABREU 00A	DLPH	$e^+e^- \rightarrow \gamma\gamma$
> 0.680	(> 0.542)	29 ABREU 00S	DLPH	$e^+e^- \rightarrow \mu^+\mu^-, \tau^+\tau^-$
> 15-28	99.7	30 CHANG 00B	RVUE	Electroweak
> 0.98	95	31 CHEUNG 00	RVUE	$e^+e^- \rightarrow \gamma\gamma$
> 0.29-0.38	95	32 GRAESSER 00	RVUE	$(g-2)_\mu$
> 0.50-1.1	95	33 HAN 00	RVUE	Electroweak
> 2.0	(> 2.0)	34 MATHEWS 00	RVUE	$p\bar{p} \rightarrow jj$
> 1.0	(> 1.1)	35 MELE 00	RVUE	$e^+e^- \rightarrow VV$
		36 ABBIENDI 99P	OPAL	
		37 ACCIARRI 99M	L3	
		38 ACCIARRI 99s	L3	
> 1.412	(> 1.077)	39 BOURILKOV 99		$e^+e^- \rightarrow e^+e^-$

- 1 KHACHATRYAN 15J use dijet angular distributions in  $19.7 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  to place a lower bound on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- 2 GIUDICE 03 place bounds on  $\Lambda_6$ , the coefficient of the gravitationally-induced dimension-6 operator  $(2\pi\lambda/\Lambda_6^2)(\sum \bar{\psi}\gamma^\mu\psi)(\sum \bar{\psi}\gamma_\mu\psi)$ , using data from a variety of experiments. Results are quoted for  $\lambda=\pm 1$  and are independent of  $\delta$ .
- 3 KHACHATRYAN 15AE use  $20.6 (19.7) \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  in the dimuon (dielectron) channel to place a lower limit on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- 4 AAD 14BE use  $20 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8 \text{ TeV}$  in the dilepton channel to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ).
- 5 AAD 13AS use  $4.9 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ).
- 6 AAD 13E use  $4.9$  and  $5.0 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  in the dielectron and dimuon channels, respectively, to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ). The dielectron and dimuon channels are combined with previous results in the diphoton channel to set the best limit. Bounds on individual channels and different priors can be found in their Table VII.
- 7 AAD 12Y use  $2.12 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ).
- 8 BAAK 12 use electroweak precision observables to place bounds on the ratio  $\Lambda_T/M_D$  as a function of  $M_D$ . See their Fig. 22 for constraints with a Higgs mass of  $120 \text{ GeV}$ .
- 9 CHATRCHYAN 12J use approximately  $2 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  in the dielectron and dimuon channels to place lower limits on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- 10 CHATRCHYAN 12R use  $2.2 \text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ).
- 11 AARON 11c search for deviations in the differential cross section of  $e^\pm p \rightarrow e^\pm X$  in  $446 \text{ pb}^{-1}$  of data taken at  $\sqrt{s} = 301$  and  $319 \text{ GeV}$  to place a bound on  $M_{TT}$ .
- 12 CHATRCHYAN 11A use  $36 \text{ pb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7 \text{ TeV}$  to place lower limits on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- 13 ABAZOV 09AE use dijet angular distributions in  $0.7 \text{ fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$  to place lower bounds on  $\Lambda_T$  (equivalent to their  $M_S$ ), here converted to  $M_{TT}$ .
- 14 ABAZOV 09D use  $1.05 \text{ fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$  to place lower bounds on  $\Lambda_T$  (equivalent to their  $M_S$ ), here converted to  $M_{TT}$ .
- 15 SCHAEEL 07A use  $e^+e^-$  collisions at  $\sqrt{s} = 189\text{--}209 \text{ GeV}$  to place lower limits on  $\Lambda_T$ , here converted to limits on  $M_{TT}$ .
- 16 ABDALLAH 06C use  $e^+e^-$  collisions at  $\sqrt{s} \sim 130\text{--}207 \text{ GeV}$  to place lower limits on  $M_{TT}$ , which is equivalent to their definition of  $M_S$ . Bound shown includes all possible final state leptons,  $\ell = e, \mu, \tau$ . Bounds on individual leptonic final states can be found in their Table 31.
- 17 GERDES 06 use  $100$  to  $110 \text{ pb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.8 \text{ TeV}$ , as recorded by the CDF Collaboration during Run I of the Tevatron. Bound shown includes a  $K$ -factor of  $1.3$ . Bounds on individual  $e^+e^-$  and  $\gamma\gamma$  final states are found in their Table I.
- 18 ABAZOV 05v use  $246 \text{ pb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$  to search for deviations in the differential cross section to  $\mu^+\mu^-$  from graviton exchange.

# Searches Particle Listings

## Extra Dimensions

- <sup>19</sup> CHEKANOV 04B search for deviations in the differential cross section of  $e^{\pm}p \rightarrow e^{\pm}X$  with  $130\text{ pb}^{-1}$  of combined data and  $Q^2$  values up to  $40,000\text{ GeV}^2$  to place a bound on  $M_{TT}$ .
- <sup>20</sup> ABBIENDI 03D use  $e^+e^-$  collisions at  $\sqrt{s}=181\text{--}209\text{ GeV}$  to place bounds on the ultraviolet scale  $M_{TT}$ , which is equivalent to their definition of  $M_S$ .
- <sup>21</sup> ACHARD 03D look for deviations in the cross section for  $e^+e^- \rightarrow ZZ$  from  $\sqrt{s}=200\text{--}209\text{ GeV}$  to place a bound on  $M_{TT}$ .
- <sup>22</sup> ADLOFF 03 search for deviations in the differential cross section of  $e^{\pm}p \rightarrow e^{\pm}X$  at  $\sqrt{s}=301$  and  $319\text{ GeV}$  to place bounds on  $M_{TT}$ .
- <sup>23</sup> GIUDICE 03 review existing experimental bounds on  $M_{TT}$  and derive a combined limit.
- <sup>24</sup> HEISTER 03C use  $e^+e^-$  collisions at  $\sqrt{s}=189\text{--}209\text{ GeV}$  to place bounds on the scale of dim-8 gravitational interactions. Their  $M_S^{\pm}$  is equivalent to our  $M_{TT}$  with  $\lambda=\pm 1$ .
- <sup>25</sup> ACHARD 02 search for s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma$  at  $E_{\text{cm}}=192\text{--}209\text{ GeV}$ .
- <sup>26</sup> ABBOTT 01 search for variations in differential cross sections to  $e^+e^-$  and  $\gamma\gamma$  final states at the Tevatron.
- <sup>27</sup> ABBIENDI 00R uses  $e^+e^-$  collisions at  $\sqrt{s}=189\text{ GeV}$ .
- <sup>28</sup> ABREU 00A search for s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma$  at  $E_{\text{cm}}=189\text{--}202\text{ GeV}$ .
- <sup>29</sup> ABREU 00s uses  $e^+e^-$  collisions at  $\sqrt{s}=183$  and  $189\text{ GeV}$ . Bounds on  $\mu$  and  $\tau$  individual final states given in paper.
- <sup>30</sup> CHANG 00B derive  $3\sigma$  limit on  $M_{TT}$  of  $(28,19,15)\text{ TeV}$  for  $\delta=(2,4,6)$  respectively assuming the presence of a torsional coupling in the gravitational action. Highly model dependent.
- <sup>31</sup> CHEUNG 00 obtains limits from anomalous diphoton production at OPAL due to graviton exchange. Original limit for  $\delta=4$ . However, unknown  $UV$  theory renders  $\delta$  dependence unreliable. Original paper works in HLZ convention.
- <sup>32</sup> GRAESSER 00 obtains a bound from graviton contributions to  $g\text{--}2$  of the muon through loops of  $0.29\text{ TeV}$  for  $\delta=2$  and  $0.38\text{ TeV}$  for  $\delta=4,6$ . Limits scale as  $\lambda^{1/2}$ . However calculational scheme not well-defined without specification of high-scale theory. See the "Extra Dimensions Review."
- <sup>33</sup> HAN 00 calculates corrections to gauge boson self-energies from KK graviton loops and constrain them using  $S$  and  $T$ . Bounds on  $M_{TT}$  range from  $0.5\text{ TeV}$  ( $\delta=6$ ) to  $1.1\text{ TeV}$  ( $\delta=2$ ); see text. Limits have strong dependence,  $\lambda^{\delta+2}$ , on unknown  $\lambda$  coefficient.
- <sup>34</sup> MATHEWS 00 search for evidence of graviton exchange in CDF and DØ dijet production data. See their Table 2 for slightly stronger  $\delta$ -dependent bounds. Limits expressed in terms of  $\tilde{M}_S^4 = M_{TT}^4/8$ .
- <sup>35</sup> MELE 00 obtains bound from KK graviton contributions to  $e^+e^- \rightarrow VV$  ( $V=\gamma, W, Z$ ) at LEP. Authors use Hewett conventions.
- <sup>36</sup> ABBIENDI 99P search for s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma$  at  $E_{\text{cm}}=189\text{ GeV}$ . The limits  $G_{\pm} > 660\text{ GeV}$  and  $G_{\pm} > 634\text{ GeV}$  are obtained from combined  $E_{\text{cm}}=183$  and  $189\text{ GeV}$  data, where  $G_{\pm}$  is a scale related to the fundamental gravity scale.
- <sup>37</sup> ACCIARRI 99M search for the reaction  $e^+e^- \rightarrow \gamma G$  and s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma, W^+W^-, ZZ, e^+e^-, \mu^+\mu^-, \tau^+\tau^-, q\bar{q}$  at  $E_{\text{cm}}=183\text{ GeV}$ . Limits on the gravity scale are listed in their Tables 1 and 2.
- <sup>38</sup> ACCIARRI 99s search for the reaction  $e^+e^- \rightarrow ZG$  and s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma, W^+W^-, ZZ, e^+e^-, \mu^+\mu^-, \tau^+\tau^-, q\bar{q}$  at  $E_{\text{cm}}=189\text{ GeV}$ . Limits on the gravity scale are listed in their Tables 1 and 2.
- <sup>39</sup> BOURILKOV 99 performs global analysis of LEP data on  $e^+e^-$  collisions at  $\sqrt{s}=183$  and  $189\text{ GeV}$ . Bound is on  $\Lambda_T$ .

### Limits on $1/R = M_c$

This section includes limits on  $1/R = M_c$ , the compactification scale in models with one TeV-sized extra dimension, due to exchange of Standard Model KK excitations. Bounds assume fermions are not in the bulk, unless stated otherwise. See the "Extra Dimensions" review for discussion of model dependence.

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;4.16</b>	95	<sup>1</sup> AAD	12CC ATLS	$pp \rightarrow \ell\bar{\ell}$
<b>&gt;6.1</b>		<sup>2</sup> BARBIERI	04 RVUE	Electroweak
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>3.8	95	<sup>3</sup> ACCOMANDO 15	RVUE	Electroweak
>3.40	95	<sup>4</sup> KHACHATRYAN 15T	CMS	$pp \rightarrow \ell X$
	95	<sup>5</sup> CHATRCHYAN 13AQ	CMS	$pp \rightarrow \ell X$
>1.38	95	<sup>6</sup> CHATRCHYAN 13W	CMS	$pp \rightarrow \gamma\gamma, \delta=6, M_D=5\text{ TeV}$
>0.715	95	<sup>7</sup> EDELHAUSER 13	RVUE	$pp \rightarrow \ell\bar{\ell} + X$
>1.40	95	<sup>8</sup> AAD	12CP ATLS	$pp \rightarrow \gamma\gamma, \delta=6, M_D=5\text{ TeV}$
>1.23	95	<sup>9</sup> AAD	12X ATLS	$pp \rightarrow \gamma\gamma, \delta=6, M_D=5\text{ TeV}$
>0.26	95	<sup>10</sup> ABAZOV	12M D0	$p\bar{p} \rightarrow \mu\mu$
>0.75	95	<sup>11</sup> BAAK	12 RVUE	Electroweak
		<sup>12</sup> FLACKE	12 RVUE	Electroweak
>0.43	95	<sup>13</sup> NISHIWAKI	12 RVUE	$H \rightarrow WW, \gamma\gamma$
>0.729	95	<sup>14</sup> AAD	11F ATLS	$pp \rightarrow \gamma\gamma, \delta=6, M_D=5\text{ TeV}$
>0.961	95	<sup>15</sup> AAD	11X ATLS	$pp \rightarrow \gamma\gamma, \delta=6, M_D=5\text{ TeV}$
>0.477	95	<sup>16</sup> ABAZOV	10P D0	$p\bar{p} \rightarrow \gamma\gamma, \delta=6, M_D=5\text{ TeV}$
>1.59	95	<sup>17</sup> ABAZOV	09AE D0	$p\bar{p} \rightarrow$ dijet, angular dist.
>0.6	95	<sup>18</sup> HAISCH	07 RVUE	$\bar{B} \rightarrow X_s\gamma$
>0.6	90	<sup>19</sup> GOGOLADZE	06 RVUE	Electroweak
>3.3	95	<sup>20</sup> CORNET	00 RVUE	Electroweak
>3.3–3.8	95	<sup>21</sup> RIZZO	00 RVUE	Electroweak

- <sup>1</sup> AAD 12CC use  $4.9$  and  $5.0\text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK  $Z/\gamma$  boson (equivalent to  $1/R = M_c$ ). The limit quoted here assumes a flat prior corresponding to when the pure  $Z/\gamma$  KK cross section term dominates. See their Section 15 for more details.
- <sup>2</sup> BARBIERI 04 use electroweak precision observables to place a lower bound on the compactification scale  $1/R$ . Both the gauge bosons and the Higgs boson are assumed to propagate in the bulk.
- <sup>3</sup> ACCOMANDO 15 use electroweak precision observables to place a lower bound on the compactification scale  $1/R$ . See their Fig. 2 for the bound as a function of  $\sin\beta$ , which parametrizes the VEV contribution from brane and bulk Higgs fields. The quoted value is for the minimum bound which occurs at  $\sin\beta = 0.45$ .
- <sup>4</sup> KHACHATRYAN 15T use  $19.7\text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s}=8\text{ TeV}$  to place a lower bound on the compactification scale  $1/R$ .
- <sup>5</sup> CHATRCHYAN 13AQ use  $5.0\text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  and a further  $3.7\text{ fb}^{-1}$  of data at  $\sqrt{s}=8\text{ TeV}$  to place a lower bound on the compactification scale  $1/R$ , in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 5 for the bound as a function of the universal bulk fermion mass parameter  $\mu$ .
- <sup>6</sup> CHATRCHYAN 13W use diphoton events with large missing transverse momentum in  $4.93\text{ fb}^{-1}$  of data produced from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c = 20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G\gamma$  occurs with an appreciable branching fraction.
- <sup>7</sup> EDELHAUSER 13 use  $19.6$  and  $20.6\text{ fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s}=8\text{ TeV}$  analyzed by the CMS Collaboration in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the second lightest Kaluza-Klein  $Z/\gamma$  boson (converted to a limit on  $1/R = M_c$ ). The bound assumes Standard Model fields propagating in the bulk and that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c = 20$ .
- <sup>8</sup> AAD 12CP use diphoton events with large missing transverse momentum in  $4.8\text{ fb}^{-1}$  of data produced from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c = 20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G\gamma$  occurs with an appreciable branching fraction.
- <sup>9</sup> AAD 12X use diphoton events with large missing transverse momentum in  $1.07\text{ fb}^{-1}$  of data produced from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c = 20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G\gamma$  occurs with an appreciable branching fraction.
- <sup>10</sup> ABAZOV 12M use same-sign dimuon events in  $7.3\text{ fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96\text{ TeV}$  to place a lower bound on the compactification scale  $1/R$ , in models with universal extra dimensions where all Standard Model fields propagate in the bulk.
- <sup>11</sup> BAAK 12 use electroweak precision observables to place a lower bound on the compactification scale  $1/R$ , in models with universal extra dimensions and Standard Model fields propagating in the bulk. Bound assumes a  $125\text{ GeV}$  Higgs mass. See their Fig. 25 for the bound as a function of the Higgs mass.
- <sup>12</sup> FLACKE 12 use electroweak precision observables to place a lower bound on the compactification scale  $1/R$ , in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 1 for the bound as a function of the universal bulk fermion mass parameter  $\mu$ .
- <sup>13</sup> NISHIWAKI 12 use up to  $2\text{ fb}^{-1}$  of data from the ATLAS and CMS experiments that constrains the production cross section of a Higgs-like particle to place a lower bound on the compactification scale  $1/R$  in universal extra dimension models. The quoted bound assumes Standard Model fields propagating in the bulk and a  $125\text{ GeV}$  Higgs mass. See their Fig. 1 for the bound as a function of the Higgs mass.
- <sup>14</sup> AAD 11F use diphoton events with large missing transverse energy in  $3.1\text{ pb}^{-1}$  of data produced from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c = 20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G\gamma$  occurs with an appreciable branching fraction.
- <sup>15</sup> AAD 11X use diphoton events with large missing transverse energy in  $36\text{ pb}^{-1}$  of data produced from  $pp$  collisions at  $\sqrt{s}=7\text{ TeV}$  to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c = 20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G\gamma$  occurs with an appreciable branching fraction.
- <sup>16</sup> ABAZOV 10P use diphoton events with large missing transverse energy in  $6.3\text{ fb}^{-1}$  of data produced from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96\text{ TeV}$  to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c=20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G\gamma$  occurs with an appreciable branching fraction.
- <sup>17</sup> ABAZOV 09AE use dijet angular distributions in  $0.7\text{ fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96\text{ TeV}$  to place a lower bound on the compactification scale.
- <sup>18</sup> HAISCH 07 use inclusive  $\bar{B}$ -meson decays to place a Higgs mass independent bound on the compactification scale  $1/R$  in the minimal universal extra dimension model.
- <sup>19</sup> GOGOLADZE 06 use electroweak precision observables to place a lower bound on the compactification scale in models with universal extra dimensions. Bound assumes a  $115\text{ GeV}$  Higgs mass. See their Fig. 3 for the bound as a function of the Higgs mass.
- <sup>20</sup> CORNET 00 translates a bound on the coefficient of the 4-fermion operator  $(\bar{\ell}\gamma_{\mu}\tau^a\ell)(\bar{\ell}\gamma^{\mu}\tau^a\ell)$  derived by Hagiwara and Matsumoto into a limit on the mass scale of KK  $W$  bosons.
- <sup>21</sup> RIZZO 00 obtains limits from global electroweak fits in models with a Higgs in the bulk ( $3.8\text{ TeV}$ ) or on the standard brane ( $3.3\text{ TeV}$ ).

### Limits on Kaluza-Klein Gravitons in Warped Extra Dimensions

This section places limits on the mass of the first Kaluza-Klein (KK) excitation of the graviton in the warped extra dimension model of Randall and Sundrum. Bounds in parenthesis assume Standard Model fields propagate in the bulk. Experimental bounds



depend strongly on the warp parameter,  $k$ . See the “Extra Dimensions” review for a full discussion.

Here we list limits for the value of the warp parameter  $k/\overline{M}_P = 0.1$ .

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;2.73</b>	95	<sup>1</sup> KHACHATRYAN15AE CMS	$pp \rightarrow e^+e^-, \mu^+\mu^-$	• • • We do not use the following data for averages, fits, limits, etc. • • •
>2.66	95	<sup>2</sup> AAD 15AD ATLS	$pp \rightarrow G \rightarrow \gamma\gamma$	
	95	<sup>3</sup> AAD 15BK ATLS	$pp \rightarrow G \rightarrow hh$	
	95	<sup>4</sup> KHACHATRYAN15R CMS	$pp \rightarrow G \rightarrow hh$	
>2.68	95	<sup>5</sup> AAD 14V ATLS	$pp \rightarrow G \rightarrow e^+e^-, \mu^+\mu^-$	
	95	<sup>6</sup> KHACHATRYAN14A CMS	$pp \rightarrow G \rightarrow WW, ZZ, WZ$	
>1.23 (> 0.84)	95	<sup>7</sup> AAD 13A ATLS	$pp \rightarrow G \rightarrow WW$	
>2.23	95	<sup>8</sup> AAD 13AS ATLS	$pp \rightarrow \gamma\gamma, e^+e^-, \mu^+\mu^-$	
>2.39	95	<sup>9</sup> CHATRCHYAN13AF CMS	$pp \rightarrow e^+e^-, \mu^+\mu^-$	
	95	<sup>10</sup> CHATRCHYAN13U CMS	$pp \rightarrow G \rightarrow ZZ$	
>0.845	95	<sup>11</sup> AAD 12AD ATLS	$pp \rightarrow G \rightarrow ZZ$	
>2.16	95	<sup>12</sup> AAD 12CC ATLS	$pp \rightarrow G \rightarrow \ell\ell$	
>1.95	95	<sup>13</sup> AAD 12Y ATLS	$pp \rightarrow \gamma\gamma, e^+e^-, \mu^+\mu^-$	
	95	<sup>14</sup> AALTONEN 12V CDF	$p\bar{p} \rightarrow G \rightarrow ZZ$	
	95	<sup>15</sup> BAAK 12 RVUE	Electroweak	
>1.84	95	<sup>16</sup> CHATRCHYAN12R CMS	$pp \rightarrow G \rightarrow \gamma\gamma$	
>1.63	95	<sup>17</sup> AAD 11AD ATLS	$pp \rightarrow G \rightarrow \ell\ell$	
	95	<sup>18</sup> AALTONEN 11G CDF	$p\bar{p} \rightarrow G \rightarrow ZZ$	
>1.058	95	<sup>19</sup> AALTONEN 11R CDF	$p\bar{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$	
>0.754	95	<sup>20</sup> ABAZOV 11H D0	$p\bar{p} \rightarrow G \rightarrow WW$	
>1.079	95	<sup>21</sup> CHATRCHYAN11 CMS	$pp \rightarrow G \rightarrow \ell\ell$	
>0.607	95	<sup>22</sup> AALTONEN 10N CDF	$p\bar{p} \rightarrow G \rightarrow WW$	
>1.05	95	<sup>23</sup> ABAZOV 10F D0	$p\bar{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$	
	95	<sup>24</sup> AALTONEN 08S CDF	$p\bar{p} \rightarrow G \rightarrow ZZ$	
>0.90	95	<sup>25</sup> ABAZOV 08J D0	$p\bar{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$	
	95	<sup>26</sup> AALTONEN 07G CDF	$p\bar{p} \rightarrow G \rightarrow \gamma\gamma$	
>0.889	95	<sup>27</sup> AALTONEN 07H CDF	$p\bar{p} \rightarrow G \rightarrow e\bar{e}$	
>0.785	95	<sup>28</sup> ABAZOV 05N D0	$p\bar{p} \rightarrow G \rightarrow \ell\ell, \gamma\gamma$	
>0.71	95	<sup>29</sup> ABULENCIA 05A CDF	$p\bar{p} \rightarrow G \rightarrow \ell\ell$	

- <sup>1</sup> KHACHATRYAN 15AE use 20.6 (19.7)  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV in the dimuon (dielectron) channel to place a lower bound on the mass of the lightest KK graviton.
- <sup>2</sup> AAD 15AD use 20.3  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV in the diphoton channel to place a lower limit on the mass of the lightest KK graviton. See their Table IV for limits with warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.
- <sup>3</sup> AAD 15BK use 19.5  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV to search for Higgs boson pair production in the  $b\bar{b}b\bar{b}$  final state, and exclude masses of the lightest KK graviton. See their Table 9 for the excluded mass ranges with warp parameter values  $k/\overline{M}_P = 1.0, 1.5$ , and 2.0.
- <sup>4</sup> KHACHATRYAN 15R use 17.9  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV to search for Higgs boson pair production in the  $b\bar{b}b\bar{b}$  final state, and exclude a KK graviton with mass from 380 to 830 GeV.
- <sup>5</sup> AAD 14V use 20  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV in the dielectron and dimuon channels to place a lower bound on the mass of the lightest KK graviton.
- <sup>6</sup> KHACHATRYAN 14A use 19.7  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV to search for KK gravitons in a warped extra dimension decaying to dibosons. See their Figure 9 for limits on the cross section times branching fraction as a function of the KK graviton mass.
- <sup>7</sup> AAD 13A use 4.7  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV to place a lower bound on the mass of the lightest KK graviton.
- <sup>8</sup> AAD 13AS use 4.9  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. The diphoton channel is combined with previous results in the dielectron and dimuon channels to set the best limit. See their Table 2 for warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.
- <sup>9</sup> CHATRCHYAN 13AF use 5.3 and 4.1  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV and 8 TeV, respectively, in the dielectron and dimuon channels, to place a lower bound on the mass of the lightest KK graviton.
- <sup>10</sup> CHATRCHYAN 13U use 5  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV to search for KK gravitons in a warped extra dimension decaying to  $ZZ$  dibosons. See their Figure 5 for limits on the lightest KK graviton mass as a function of  $k/\overline{M}_P$ .
- <sup>11</sup> AAD 12AD use 1.02  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV to search for KK gravitons in a warped extra dimension decaying to  $ZZ$  dibosons in the  $lljj$  and  $llll$  channels ( $\ell = e, \mu$ ). The limit is quoted for the combined  $lljj + llll$  channels. See their Figure 5 for limits on the cross section  $\sigma(G \rightarrow ZZ)$  as a function of the graviton mass.
- <sup>12</sup> AAD 12CC use 4.9 and 5.0  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK graviton. See their Figure 5 for limits on the lightest KK graviton mass as a function of  $k/\overline{M}_P$ .
- <sup>13</sup> AAD 12Y use 2.12  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. The diphoton channel is combined with previous results in the dielectron and dimuon channels to set the best limit. See their Table 3 for warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.
- <sup>14</sup> AALTONEN 12V use 6  $\text{fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in a warped extra dimension decaying to  $ZZ$  dibosons in the  $lljj$  and  $llll$  channels ( $\ell = e, \mu$ ). It provides improved limits over the previous analysis in AALTONEN 11G. See their Figure 16 for limits from all channels combined on the cross section times branching ratio  $\sigma(p\bar{p} \rightarrow G^* \rightarrow ZZ)$  as a function of the graviton mass.
- <sup>15</sup> BAAK 12 use electroweak precision observables to place a lower bound on the compactification scale  $k e^{-\pi k R}$ , assuming Standard Model fields propagate in the bulk and the Higgs is confined to the IR brane. See their Fig. 27 for more details.
- <sup>16</sup> CHATRCHYAN 12R use 2.2  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. See their Table III for warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.

- <sup>17</sup> AAD 11AD use 1.08 and 1.21  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest graviton. For warp parameter values  $k/\overline{M}_P$  between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 0.71 and 1.63 TeV. See their Table IV for more details.
- <sup>18</sup> AALTONEN 11G use 2.5–2.9  $\text{fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in a warped extra dimension decaying to  $ZZ$  dibosons via the  $eee, ee\mu\mu, \mu\mu\mu\mu, eejj$ , and  $\mu\mu jj$  channels. See their Fig. 20 for limits on the cross section  $\sigma(G \rightarrow ZZ)$  as a function of the graviton mass.
- <sup>19</sup> AALTONEN 11R uses 5.7  $\text{fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV in the dielectron channel to place a lower bound on the mass of the lightest graviton. It provides combined limits with the diphoton channel analysis of AALTONEN 11U. For warp parameter values  $k/\overline{M}_P$  between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 612 and 1058 GeV. See their Table I for more details.
- <sup>20</sup> ABAZOV 11H use 5.4  $\text{fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to place a lower bound on the mass of the lightest graviton. Their 95% C.L. exclusion limit does not include masses less than 300 GeV.
- <sup>21</sup> CHATRCHYAN 11 use 35 and 40  $\text{pb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest graviton. For a warp parameter value  $k/\overline{M}_P = 0.05$ , the lower limit on the mass of the lightest graviton is 0.855 TeV.
- <sup>22</sup> AALTONEN 10N use 2.9  $\text{fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to place a lower bound on the mass of the lightest graviton.
- <sup>23</sup> ABAZOV 10F use 5.4  $\text{fb}^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to place a lower bound on the mass of the lightest graviton. For warp parameter values of  $k/\overline{M}_P$  between 0.01 and 0.1 the lower limit on the mass of the lightest graviton is between 560 and 1050 GeV. See their Fig. 3 for more details.
- <sup>24</sup> AALTONEN 08S use  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to four electrons via two  $Z$  bosons using 1.1  $\text{fb}^{-1}$  of data. See their Fig. 8 for limits on  $\sigma \cdot \text{B}(G \rightarrow ZZ)$  versus the graviton mass.
- <sup>25</sup> ABAZOV 08J use  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons and photons using 1  $\text{fb}^{-1}$  of data. For warp parameter values of  $k/\overline{M}_P$  between 0.01 and 0.1 the lower limit on the mass of the lightest excitation is between 300 and 900 GeV. See their Fig. 4 for more details.
- <sup>26</sup> AALTONEN 07G use  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to photons using 1.2  $\text{fb}^{-1}$  of data. For warp parameter values of  $k/\overline{M}_P = 0.1, 0.05$ , and 0.01 the bounds on the graviton mass are 850, 694, and 230 GeV, respectively. See their Fig. 3 for more details. See also AALTONEN 07H.
- <sup>27</sup> AALTONEN 07H use  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons using 1.3  $\text{fb}^{-1}$  of data. For a warp parameter value of  $k/\overline{M}_P = 0.1$  the bound on the graviton mass is 807 GeV. See their Fig. 4 for more details. A combined analysis with the diphoton data of AALTONEN 07G yields for  $k/\overline{M}_P = 0.1$  a graviton mass lower bound of 889 GeV.
- <sup>28</sup> ABAZOV 05N use  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons, electrons or photons, using 260  $\text{pb}^{-1}$  of data. For warp parameter values of  $k/\overline{M}_P = 0.1, 0.05$ , and 0.01, the bounds on the graviton mass are 785, 650 and 250 GeV respectively. See their Fig. 3 for more details.
- <sup>29</sup> ABULENCIA 05A use  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons or electrons, using 200  $\text{pb}^{-1}$  of data. For warp parameter values of  $k/\overline{M}_P = 0.1, 0.05$ , and 0.01, the bounds on the graviton mass are 710, 510 and 170 GeV respectively.

### Limits on Kaluza-Klein Gluons in Warped Extra Dimensions

This section places limits on the mass of the first Kaluza-Klein (KK) excitation of the gluon in warped extra dimension models with Standard Model fields propagating in the bulk. Bounds are given for a specific benchmark model with  $\Gamma/m = 15.3\%$  where  $\Gamma$  is the width and  $m$  the mass of the KK gluon. See the “Extra Dimensions” review for more discussion.

VALUE (TeV)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;2.5</b>	95	<sup>1</sup> CHATRCHYAN13BM CMS	$g_{KK} \rightarrow t\bar{t}$	• • • We do not use the following data for averages, fits, limits, etc. • • •
>2.07	95	<sup>2</sup> AAD 13AQ ATLS	$g_{KK} \rightarrow t\bar{t} \rightarrow \ell j$	
	95	<sup>3</sup> CHEN 13A	$\bar{B} \rightarrow X_s \gamma$	
>1.5	95	<sup>4</sup> AAD 12BV ATLS	$g_{KK} \rightarrow t\bar{t} \rightarrow \ell j$	

<sup>1</sup> CHATRCHYAN 13BM use 19.7  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 8$  TeV. Bound is for a width of approximately 15–20% of the KK gluon mass.

<sup>2</sup> AAD 13AQ use 4.7  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV.

<sup>3</sup> CHEN 13A place limits on the KK mass scale for a specific warped model with custodial symmetry and bulk fermions. See their Figures 4 and 5.

<sup>4</sup> AAD 12BV use 2.05  $\text{fb}^{-1}$  of data from  $pp$  collisions at  $\sqrt{s} = 7$  TeV.

### REFERENCES FOR Extra Dimensions

AAD	15AD	PR D92 032004	G. Aad et al.	(ATLAS Collab.)
AAD	15BK	EPJ C75 412	G. Aad et al.	(ATLAS Collab.)
AAD	15CS	PR D91 012008	G. Aad et al.	(ATLAS Collab.)
Also		PR D92 059903 (err.)	G. Aad et al.	(ATLAS Collab.)
ACCOMANDO	15	MPL A30 1540010	E. Accomando	(SHMP)
KHACHATRYAN	15AE	JHEP 1504 025	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15AL	EPJ C75 235	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15J	PL B746 79	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15R	PL B749 560	V. Khachatryan et al.	(CMS Collab.)
KHACHATRYAN	15T	PR D91 092005	V. Khachatryan et al.	(CMS Collab.)
AAD	14BE	EPJ C74 3134	G. Aad et al.	(ATLAS Collab.)
AAD	14V	PR D90 052005	G. Aad et al.	(ATLAS Collab.)
KHACHATRYAN	14A	JHEP 1408 174	V. Khachatryan et al.	(CMS Collab.)
AAD	13A	PL B718 860	G. Aad et al.	(ATLAS Collab.)
AAD	13AD	JHEP 1304 075	G. Aad et al.	(ATLAS Collab.)
AAD	13AQ	PR D88 012004	G. Aad et al.	(ATLAS Collab.)
AAD	13AS	NJP 15 043007	G. Aad et al.	(ATLAS Collab.)
AAD	13C	PRL 110 011802	G. Aad et al.	(ATLAS Collab.)
AAD	13D	JHEP 1301 029	G. Aad et al.	(ATLAS Collab.)
AAD	13E	PR D87 015010	G. Aad et al.	(ATLAS Collab.)

# Searches Particle Listings

## Extra Dimensions, WIMPs and Other Particle Searches

CHATRCHYAN	13AF	PL B720 63	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	13AQ	PR D87 072005	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	13BM	PRL 111 211804	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
Also		PRL 112 119903 (errat.)	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	13U	JHEP 1302 036	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	13W	JHEP 1303 111	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHEN	13A	CPC 37 063102	J.-B. Chen <i>et al.</i>	(DALI)
EDELHAUSER	13	JHEP 1308 091	L. Edelhauser, T. Flacke, M. Kramer	(AACH, KAIST)
XU	13	JP G40 035107	J. Xu <i>et al.</i>	(ATLAS Collab.)
AAD	12AD	PL B712 331	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	12BV	JHEP 1209 041	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	12CC	JHEP 1211 138	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	12CP	PL B718 411	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	12X	PL B710 519	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	12Y	PL B710 538	G. Aad <i>et al.</i>	(ATLAS Collab.)
AALTONEN	12V	PR D85 012008	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	12M	PRL 108 131802	V.M. Abazov <i>et al.</i>	(D0 Collab.)
AJELLO	12	JCAP 1202 012	M. Ajello <i>et al.</i>	(Fermi-LAT Collab.)
BAAK	12	EPI C72 2003	M. Baak <i>et al.</i>	(Glitter Group)
CHATRCHYAN	12AP	JHEP 1209 094	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	12J	PL B711 15	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	12R	PRL 108 111801	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
FLACKE	12	PR D85 126007	T. Flacke, C. Pasold	(WURZ)
NISHIWAKI	12	PL B707 506	K. Nishiwaki <i>et al.</i>	(KOBE, OSAK)
AAD	11AD	PRL 107 272002	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	11F	PRL 106 121803	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	11S	PL B705 294	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	11X	EPJ C71 1744	G. Aad <i>et al.</i>	(ATLAS Collab.)
AALTONEN	11G	PR D83 112008	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11R	PRL 107 051801	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	11U	PR D83 011102	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AARON	11C	PL B705 52	F. D. Aaron <i>et al.</i>	(H1 Collab.)
ABAZOV	11H	PRL 107 011801	V. M. Abazov <i>et al.</i>	(D0 Collab.)
BEZERRA	11	PR D83 075004	V.B. Bezerra <i>et al.</i>	(D0 Collab.)
CHATRCHYAN	11	JHEP 1105 093	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	11A	JHEP 1105 085	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	11U	PRL 107 201804	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
SUSHKOV	11	PRL 107 171101	A.O. Sushkov <i>et al.</i>	(CDF Collab.)
AALTONEN	10N	PRL 104 241801	T. Aaltonen <i>et al.</i>	(D0 Collab.)
ABAZOV	10F	PRL 104 241802	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	10P	PRL 105 221802	V.M. Abazov <i>et al.</i>	(D0 Collab.)
BEZERRA	10	PR D81 055003	V.B. Bezerra <i>et al.</i>	(D0 Collab.)
ABAZOV	09AE	PRL 103 191803	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	09D	PRL 102 051601	V.M. Abazov <i>et al.</i>	(D0 Collab.)
MASUDA	09	PRL 102 171101	M. Masuda, M. Sasaki	(ICRR)
AALTONEN	08AC	PRL 101 181602	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	08S	PR D78 012008	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ABAZOV	08J	PRL 100 091802	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	08S	PRL 101 011601	V.M. Abazov <i>et al.</i>	(D0 Collab.)
DAS	08	PR D78 063011	P.K. Das, V.H.S. Kumar, P.K. Suresh	(STAN)
GERACI	08	PR D78 022002	A.A. Geraci <i>et al.</i>	(STAN)
TRENKEL	08	PR D77 122001	C. Trenkel	(CDF Collab.)
AALTONEN	07G	PRL 99 171801	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	07H	PRL 99 171802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
DECCA	07A	EPJ C51 963	R.S. Decca <i>et al.</i>	(CDF Collab.)
HAISCH	07	PR D76 034014	U. Haisch, A. Weiler	(ALEPH Collab.)
KAPNER	07	PRL 98 021101	D.J. Kapner <i>et al.</i>	(ALEPH Collab.)
SCHAEI	07A	EPJ C49 411	S. Schaei <i>et al.</i>	(ALEPH Collab.)
TU	07	PRL 98 201101	L.-C. Tu <i>et al.</i>	(DELPHI Collab.)
ABDALLAH	06C	EPJ C45 589	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABULENCIA,A	06	PRL 97 171802	A. Abulencia <i>et al.</i>	(CDF Collab.)
GERDES	06	PR D73 112008	D. Gerdes <i>et al.</i>	(CDF Collab.)
GOGOLADZE	06	PR D74 093012	I. Gogoladze, C. Macesanu	(D0 Collab.)
ABAZOV	05N	PRL 95 091801	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	05V	PRL 95 161602	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABDALLAH	05B	EPJ C38 395	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABULENCIA	05A	PRL 95 252001	A. Abulencia <i>et al.</i>	(CDF Collab.)
SMULLIN	05	PR D72 122001	S.J. Smullin <i>et al.</i>	(L3 Collab.)
ACHARD	04E	PL B587 16	P. Achard <i>et al.</i>	(CDF Collab.)
ACOSTA	04C	PRL 92 121802	D. Acosta <i>et al.</i>	(CDF Collab.)
BARBIERI	04	NP B703 127	R. Barbieri <i>et al.</i>	(CDF Collab.)
CASSE	04	PL 92 111102	M. Casse <i>et al.</i>	(ZEUS Collab.)
CHEKANOV	04B	PL B591 23	S. Chekanov <i>et al.</i>	(WASH)
HOYLE	04	PR D70 042004	C.D. Hoyle <i>et al.</i>	(OPAL Collab.)
ABAZOV	03	PRL 90 251802	V.M. Abazov <i>et al.</i>	(L3 Collab.)
ABBIENDI	03D	EPJ C26 331	G. Abbiendi <i>et al.</i>	(H1 Collab.)
ACHARD	03D	PL B572 133	P. Achard <i>et al.</i>	(H1 Collab.)
ADLOFF	03	PL B568 35	C. Adloff <i>et al.</i>	(H1 Collab.)
CHIAVERINI	03	PRL 90 151101	J. Chiaverini <i>et al.</i>	(H1 Collab.)
GIUDICE	03	NP B663 377	G.F. Giudice, A. Strumia	(H1 Collab.)
HANNESTAD	03	PR D67 125008	S. Hannestad, G.G. Raffelt	(H1 Collab.)
Also		PR D69 029901(errat.)	S. Hannestad, G.G. Raffelt	(H1 Collab.)
HEISTER	03C	EPJ C28 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
LONG	03	Nature 421 922	J.C. Long <i>et al.</i>	(L3 Collab.)
ACHARD	02	PL B524 65	P. Achard <i>et al.</i>	(L3 Collab.)
ACHARD	02D	PL B531 28	P. Achard <i>et al.</i>	(L3 Collab.)
HANNESTAD	02	PRL 88 071301	S. Hannestad, G. Raffelt	(D0 Collab.)
ABBOTT	01	PRL 86 1156	B. Abbott <i>et al.</i>	(D0 Collab.)
FAIRBAIRN	01	PL B508 335	M. Fairbairn	(D0 Collab.)
HANHART	01	PL B509 1	C. Hanhart <i>et al.</i>	(D0 Collab.)
HOYLE	01	PRL 86 1418	C.D. Hoyle <i>et al.</i>	(D0 Collab.)
ABBIENDI	00R	EPJ C13 553	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	00A	PL B491 67	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00S	PL B485 45	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00Z	EPJ C17 53	P. Abreu <i>et al.</i>	(DELPHI Collab.)
CASSISI	00	PL B481 323	S. Cassisi <i>et al.</i>	(DELPHI Collab.)
CHANG	00B	PRL 85 3765	L.N. Chang <i>et al.</i>	(DELPHI Collab.)
CHEUNG	00	PR D61 015005	K. Cheung	(DELPHI Collab.)
CORNET	00	PR D61 037701	F. Cornet, M. Relano, J. Rico	(DELPHI Collab.)
GRAESSER	00	PR D61 074019	M.L. Graesser	(DELPHI Collab.)
HAN	00	PR D62 125018	T. Han, D. Marfatia, R.-J. Zhang	(DELPHI Collab.)
MATTHEWS	00	JHEP 0007 008	P. Matthews, S. Raychaudhuri, K. Sridhar	(DELPHI Collab.)
MELE	00	PR D61 117901	S. Mele, E. Sanchez	(DELPHI Collab.)
RIZZO	00	PR D61 016007	T.G. Rizzo, J.D. Wells	(DELPHI Collab.)
ABBIENDI	99P	PL B465 303	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ACCIARRI	99M	PL B464 135	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	99R	PL B470 268	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	99S	PL B470 281	M. Acciari <i>et al.</i>	(L3 Collab.)
BOURLIKOV	99	JHEP 9908 006	D. Bourlikov	(L3 Collab.)
HOSKINS	85	PR D32 3084	J.K. Hoskins <i>et al.</i>	(L3 Collab.)

## WIMPs and Other Particles Searches for

### OMITTED FROM SUMMARY TABLE WIMPS AND OTHER PARTICLE SEARCHES

Revised August 2013 by K. Hikasa (Tohoku University).

We collect here those searches which do not appear in any of the above search categories. These are listed in the following order:

1. Galactic WIMP (weakly-interacting massive particle) searches
2. Concentration of stable particles in matter
3. General new physics searches
4. Limits on jet-jet resonance in hadron collisions
5. Limits on neutral particle production at accelerators
6. Limits on charged particles in  $e^+e^-$  collisions
7. Limits on charged particles in hadron reactions
8. Limits on charged particles in cosmic rays
9. Searches for quantum black hole production

Note that searches appear in separate sections elsewhere for Higgs bosons (and technipions), other heavy bosons (including  $W_R$ ,  $W'$ ,  $Z'$ , leptoquarks, axiglons), axions (including pseudo-Goldstone bosons, Majorons, familons), heavy leptons, heavy neutrinos, free quarks, monopoles, supersymmetric particles, and compositeness. We include specific WIMP searches in the appropriate sections when they yield limits on hypothetical particles such as supersymmetric particles, axions, massive neutrinos, monopoles, *etc.*

We omit papers on CHAMP's, millicharged particles, and other exotic particles. We no longer list for limits on tachyons and centauros. See our 1994 edition for these limits.

### GALACTIC WIMP SEARCHES

These limits are for weakly-interacting stable particles that may constitute the invisible mass in the galaxy. Unless otherwise noted, a local mass density of  $0.3 \text{ GeV}/\text{cm}^3$  is assumed; see each paper for velocity distribution assumptions. In the papers the limit is given as a function of the  $X^0$  mass. Here we list limits only for typical mass values of 20 GeV, 100 GeV, and 1 TeV. Specific limits on supersymmetric dark matter particles may be found in the Supersymmetry section.

### Limits for Spin-Independent Cross Section of Dark Matter Particle ( $X^0$ ) on Nucleon

Isoscalar coupling is assumed to extract the limits from those on  $X^0$ -nuclei cross section.

### For $m_{X^0} = 20 \text{ GeV}$

For limits from  $X^0$  annihilation in the Sun, the assumed annihilation final state is shown in parenthesis in the comment.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • •		We do not use the following data for averages, fits, limits, etc. • • •		
$<7.3 \times 10^{-7}$	90	AGNES	16 D550	Ar
$<1 \times 10^{-5}$	90	AGNES	15 DSID	Ar
$<1.5 \times 10^{-6}$	90	<sup>1</sup> AGNESE	15A CDM2	Ge
$<1.5 \times 10^{-7}$	90	<sup>2</sup> AGNESE	15B CDM2	Ge
$<2 \times 10^{-6}$	90	<sup>3</sup> AMOLE	15 PICO	C <sub>3</sub> F <sub>8</sub>
$<1.2 \times 10^{-5}$	90	CHOI	15 SKAM	H, solar $\nu$ ( $b\bar{b}$ )
$<1.19 \times 10^{-6}$	90	CHOI	15 SKAM	H, solar $\nu$ ( $\tau^+ \tau^-$ )
$<2 \times 10^{-8}$	90	<sup>4</sup> XIAO	15 PANX	Xe
$<2.0 \times 10^{-7}$	90	<sup>5</sup> AGNESE	14 SCDM	Ge
$<3.7 \times 10^{-5}$	90	<sup>6</sup> AGNESE	14A SCDM	Ge
$<1 \times 10^{-9}$	90	<sup>7</sup> AKERIB	14 LUX	Xe
$<2 \times 10^{-6}$	90	<sup>8</sup> ANGLOHER	14 CRES	CaWO <sub>4</sub>
$<5 \times 10^{-6}$	90	FELIZARDO	14 SMP L	C <sub>2</sub> ClF <sub>5</sub>
$<8 \times 10^{-6}$	90	<sup>9</sup> LEE	14A KIMS	Csl
$<2 \times 10^{-4}$	90	<sup>10</sup> LIU	14A CDEX	Ge

See key on page 601

# Searches Particle Listings

## WIMPs and Other Particle Searches

$<1 \times 10^{-5}$	90	11	YUE	14	CDEX	Ge
$<1.08 \times 10^{-4}$	90	12	AARTSEN	13	ICCB	H, solar $\nu$ ( $\tau^+ \tau^-$ )
$<1.5 \times 10^{-5}$	90	13	ABE	13B	XMAS	Xe
$<3.1 \times 10^{-6}$	90	14	AGNESE	13	CDM2	Si
$<3.4 \times 10^{-6}$	90	15	AGNESE	13A	CDM2	Si
$<2.2 \times 10^{-6}$	90	16	AGNESE	13A	CDM2	Si
$<5 \times 10^{-5}$	90	17	LI	13B	TEXO	Ge
		18	ZHAO	13	CDEX	Ge
$<1.2 \times 10^{-7}$	90		AKIMOV	12	ZEP3	Xe
		19	ANGLOHER	12	CRES	CaWO <sub>4</sub>
$<8 \times 10^{-6}$	90	20	ANGLOHER	12	CRES	CaWO <sub>4</sub>
$<7 \times 10^{-9}$	90	21	APRILE	12	X100	Xe
		22	ARCHAMBAUD	12	PICA	F (C <sub>4</sub> F <sub>10</sub> )
$<7 \times 10^{-7}$	90	23	ARMENGAUD	12	EDE2	Ge
		24	BARRETO	12	DMIC	CCD
$<2 \times 10^{-6}$	90		BEHNKE	12	COUP	CF <sub>3</sub> I
$<7 \times 10^{-6}$	90	25	FELIZARDO	12	SMPL	C <sub>2</sub> ClF <sub>5</sub>
$<1.5 \times 10^{-6}$	90		KIM	12	KIMS	CsI
$<5 \times 10^{-5}$	90	26	AALSETH	11	CGNT	Ge
		27	AALSETH	11A	CGNT	Ge
$<5 \times 10^{-7}$	90	28	AHMED	11	CDM2	Ge, inelastic
$<2.7 \times 10^{-7}$	90	29	AHMED	11A	RVUE	Ge
		30	AHMED	11B	CDM2	Ge, low threshold
$<3 \times 10^{-6}$	90	31	ANGLE	11	XE10	Xe
$<7 \times 10^{-8}$	90	32	APRILE	11	X100	Xe
		33	APRILE	11A	X100	Xe, inelastic
$<2 \times 10^{-8}$	90	21	APRILE	11B	X100	Xe
		34	HORN	11	ZEP3	Xe
$<2 \times 10^{-7}$	90		AHMED	10	CDM2	Ge
$<1 \times 10^{-5}$	90	35	AKERIB	10	CDM2	Si, Ge, low threshold
$<1 \times 10^{-7}$	90		APRILE	10	X100	Xe
$<2 \times 10^{-6}$	90		ARMENGAUD	10	EDE2	Ge
$<4 \times 10^{-5}$	90		FELIZARDO	10	SMPL	C <sub>2</sub> ClF <sub>3</sub>
$<1.5 \times 10^{-7}$	90	36	AHMED	09	CDM2	Ge
$<2 \times 10^{-4}$	90	37	LIN	09	TEXO	Ge
		38	AALSETH	08	CGNT	Ge

- 1 AGNESE 15A reanalyse AHMED 11B low threshold data. See their Fig. 12 (left) for improved limits extending down to 5 GeV.
- 2 AGNESE 15B reanalyse AHMED 10 data.
- 3 See their Fig. 7 for limits extending down to 4 GeV.
- 4 See their Fig. 13 for limits extending down to 5 GeV.
- 5 This limit value is provided by the authors. See their Fig. 4 for limits extending down to  $m_{\chi^0} = 3.5$  GeV.
- 6 This limit value is provided by the authors. AGNESE 14A result is from CDMSlite mode operation with enhanced sensitivity to low mass  $m_{\chi^0}$ . See their Fig. 3 for limits extending down to  $m_{\chi^0} = 3.5$  GeV (see also Fig. 4 in AGNESE 14).
- 7 See their Fig. 5 for limits extending down to  $m_{\chi^0} = 5.5$  GeV.
- 8 See their Fig. 5 for limits extending down to  $m_{\chi^0} = 1$  GeV.
- 9 See their Fig. 5 for limits extending down to  $m_{\chi^0} = 5$  GeV.
- 10 LIU 14A result is based on prototype CDEX-0 detector. See their Fig. 13 for limits extending down to  $m_{\chi^0} = 2$  GeV.
- 11 See their Fig. 4 for limits extending down to  $m_{\chi^0} = 4.5$  GeV.
- 12 AARTSEN 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between June 2010 and May 2011.
- 13 See their Fig. 8 for limits extending down to  $m_{\chi^0} = 7$  GeV.
- 14 This limit value is provided by the authors. AGNESE 13 use data taken between Oct. 2006 and July 2007. See their Fig. 4 for limits extending down to  $m_{\chi^0} = 7$  GeV.
- 15 This limit value is provided by the authors. AGNESE 13A use data taken between July 2007 and Sep. 2008. Three candidate events are seen. Assuming these events are real, the best fit parameters are  $m_{\chi^0} = 8.6$  GeV and  $\sigma = 1.9 \times 10^{-5}$  pb.
- 16 This limit value is provided by the authors. Limit from combined data of AGNESE 13 and AGNESE 13A. See their Fig. 4 for limits extending down to  $m_{\chi^0} = 5.5$  GeV.
- 17 See their Fig. 4 for limits extending down to  $m_{\chi^0} = 4$  GeV.
- 18 See their Fig. 5 for limits for  $m_{\chi^0} = 4$ –12 GeV.
- 19 ANGLOHER 12 observe excess events above the expected background which are consistent with  $X^0$  with mass  $\sim 25$  GeV (or 12 GeV) and spin-independent  $X^0$ -nucleon cross section of  $2 \times 10^{-6}$  pb (or  $4 \times 10^{-5}$  pb).
- 20 Reanalysis of ANGLOHER 09 data with all three nuclides. See also BROWN 12.
- 21 See also APRILE 14A.
- 22 See their Fig. 7 for cross section limits for  $m_{\chi^0}$  between 4 and 12 GeV.
- 23 See their Fig. 4 for limits extending down to  $m_{\chi^0} = 7$  GeV.
- 24 See their Fig. 13 for cross section limits for  $m_{\chi^0}$  between 1.2 and 10 GeV.
- 25 See also DAHL 12 for a criticism.
- 26 See their Fig. 4 for limits extending to  $m_{\chi^0} = 3.5$  GeV.
- 27 AALSETH 11A find indications of annual modulation of the data, the energy spectrum being compatible with  $X^0$  mass around 8 GeV. See also AALSETH 13.
- 28 AHMED 11 search for  $X^0$  inelastic scattering. See their Fig. 8–10 for limits. The inelastic cross section reduces to the elastic cross section at the limit of zero mass splitting (Fig. 8, left).
- 29 AHMED 11A combine CDMS II and EDELWEISS data.
- 30 AHMED 11B give limits on spin-independent  $X^0$ -nucleon cross section for  $m_{\chi^0} = 4$ –12 GeV in the range  $10^{-3}$ – $10^{-5}$  pb. See their Fig. 3.
- 31 See their Fig. 3 for limits down to  $m_{\chi^0} = 4$  GeV.

- 32 APRILE 11 reanalyze APRILE 10 data.
- 33 APRILE 11A search for  $X^0$  inelastic scattering. See their Fig. 2 and 3 for limits. See also APRILE 14A.
- 34 HORN 11 perform detector calibration by neutrons. Earlier results are only marginally affected.
- 35 See their Fig. 10 and 12 for limits extending to  $X^0$  mass of 1 GeV.
- 36 Superseded by AHMED 10.
- 37 See their Fig. 6(a) for cross section limits for  $m_{\chi^0}$  extending down to 2 GeV.
- 38 See their Fig. 2 for cross section limits for  $m_{\chi^0}$  between 4 and 10 GeV.

### For $m_{\chi^0} = 100$ GeV

For limits from  $X^0$  annihilation in the Sun, the assumed annihilation final state is shown in parenthesis in the comment.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.0 \times 10^{-8}$	90	AGNES	16	D550 Ar
$<6 \times 10^{-8}$	90	AGNES	15	DSID Ar
$<4 \times 10^{-8}$	90	1 AGNESE	15B	CDM2 Ge
$<7.13 \times 10^{-6}$	90	CHOI	15	SKAM H, solar $\nu$ ( $b\bar{b}$ )
$<6.26 \times 10^{-7}$	90	CHOI	15	SKAM H, solar $\nu$ ( $W^+ W^-$ )
$<2.76 \times 10^{-7}$	90	CHOI	15	SKAM H, solar $\nu$ ( $\tau^+ \tau^-$ )
$<1.5 \times 10^{-8}$	90	XIAO	15	PANX Xe
$<1 \times 10^{-9}$	90	AKERIB	14	LUX Xe
$<4.0 \times 10^{-6}$	90	2 AVRORIN	14	BAIK H, solar $\nu$ ( $W^+ W^-$ )
$<1.0 \times 10^{-4}$	90	2 AVRORIN	14	BAIK H, solar $\nu$ ( $b\bar{b}$ )
$<1.6 \times 10^{-6}$	90	2 AVRORIN	14	BAIK H, solar $\nu$ ( $\tau^+ \tau^-$ )
$<5 \times 10^{-6}$	90	FELIZARDO	14	SMPL C <sub>2</sub> ClF <sub>5</sub>
$<6.01 \times 10^{-7}$	90	3 AARTSEN	13	ICCB H, solar $\nu$ ( $W^+ W^-$ )
$<3.30 \times 10^{-5}$	90	3 AARTSEN	13	ICCB H, solar $\nu$ ( $b\bar{b}$ )
$<1.9 \times 10^{-6}$	90	4 ADRIAN-MAR.	13	ANTR H, solar $\nu$ ( $W^+ W^-$ )
$<1.2 \times 10^{-4}$	90	4 ADRIAN-MAR.	13	ANTR H, solar $\nu$ ( $b\bar{b}$ )
$<7.6 \times 10^{-7}$	90	4 ADRIAN-MAR.	13	ANTR H, solar $\nu$ ( $\tau^+ \tau^-$ )
$<2 \times 10^{-6}$	90	5 AGNESE	13	CDM2 Si
$<1.6 \times 10^{-6}$	90	6 BOLIEV	13	BAKS H, solar $\nu$ ( $W^+ W^-$ )
$<1.9 \times 10^{-5}$	90	6 BOLIEV	13	BAKS H, solar $\nu$ ( $b\bar{b}$ )
$<7.1 \times 10^{-7}$	90	6 BOLIEV	13	BAKS H, solar $\nu$ ( $\tau^+ \tau^-$ )
$<1.67 \times 10^{-6}$	90	7 ABBASI	12	ICCB H, solar $\nu$ ( $W^+ W^-$ )
$<1.07 \times 10^{-4}$	90	7 ABBASI	12	ICCB H, solar $\nu$ ( $b\bar{b}$ )
$<4 \times 10^{-8}$	90	AKIMOV	12	ZEP3 Xe
$<1.4 \times 10^{-6}$	90	8 ANGLOHER	12	CRES CaWO <sub>4</sub>
$<3 \times 10^{-9}$	90	9 APRILE	12	X100 Xe
$<3 \times 10^{-7}$	90	BEHNKE	12	COUP CF <sub>3</sub> I
$<7 \times 10^{-6}$	90	FELIZARDO	12	SMPL C <sub>2</sub> ClF <sub>5</sub>
$<2.5 \times 10^{-7}$	90	10 KIM	12	KIMS CsI
$<2 \times 10^{-4}$	90	AALSETH	11	CGNT Ge
		11 AHMED	11	CDM2 Ge, inelastic
		12 AHMED	11A	RVUE Ge
		13 AJELLO	11	FLAT
$<3 \times 10^{-8}$	90	14 APRILE	11	X100 Xe
		15 APRILE	11A	X100 Xe, inelastic
$<1 \times 10^{-8}$	90	9 APRILE	11B	X100 Xe
$<5 \times 10^{-8}$	90	16 ARMENGAUD	11	EDE2 Ge
		17 HORN	11	ZEP3 Xe
$<4 \times 10^{-8}$	90	AHMED	10	CDM2 Ge
$<9 \times 10^{-6}$	90	AKERIB	10	CDM2 Si, Ge, low threshold
		18 AKIMOV	10	ZEP3 Xe, inelastic
$<5 \times 10^{-8}$	90	APRILE	10	X100 Xe
$<1 \times 10^{-7}$	90	ARMENGAUD	10	EDE2 Ge
$<3 \times 10^{-5}$	90	FELIZARDO	10	SMPL C <sub>2</sub> ClF <sub>3</sub>
$<5 \times 10^{-8}$	90	19 AHMED	09	CDM2 Ge
		20 ANGLE	09	XE10 Xe, inelastic
$<3 \times 10^{-4}$	90	LIN	09	TEXO Ge
		21 GIULIANI	05	RVUE

- 1 AGNESE 15B reanalyse AHMED 10 data.
- 2 AVRORIN 14 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun in data taken between 1998 and 2003. See their Table 1 for limits assuming annihilation into neutrino pairs.
- 3 AARTSEN 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between June 2010 and May 2011.
- 4 ADRIAN-MARTINEZ 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between Jan. 2007 and Dec. 2008.
- 5 AGNESE 13 use data taken between Oct. 2006 and July 2007.
- 6 BOLIEV 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken from 1978 to 2009. See also SUVOROVA 13 for an older analysis of the same data.
- 7 ABBASI 12 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.
- 8 Reanalysis of ANGLOHER 09 data with all three nuclides. See also BROWN 12.
- 9 See also APRILE 14A.
- 10 See their Fig. 6 for a limit on inelastically scattering  $X^0$  for  $m_{\chi^0} = 70$  GeV.
- 11 AHMED 11 search for  $X^0$  inelastic scattering. See their Fig. 8–10 for limits.
- 12 AHMED 11A combine CDMS and EDELWEISS data.
- 13 AJELLO 11 search for  $e^\pm$  flux from  $X^0$  annihilations in the Sun. Models in which  $X^0$  annihilates into an intermediate long-lived weakly interacting particles or  $X^0$  scatters inelastically are constrained. See their Fig. 6–8 for limits.
- 14 APRILE 11 reanalyze APRILE 10 data.

# Searches Particle Listings

## WIMPs and Other Particle Searches

- <sup>15</sup> APRILE 11A search for  $X^0$  inelastic scattering. See their Fig. 2 and 3 for limits. See also APRILE 14A.
- <sup>16</sup> Supersedes ARMENGAUD 10. A limit on inelastic cross section is also given.
- <sup>17</sup> HORN 11 perform detector calibration by neutrons. Earlier results are only marginally affected.
- <sup>18</sup> AKIMOV 10 give cross section limits for inelastically scattering dark matter. See their Fig. 4.
- <sup>19</sup> Superseded by AHMED 10.
- <sup>20</sup> ANGLE 09 search for  $X^0$  inelastic scattering. See their Fig. 4 for limits.
- <sup>21</sup> GIULIANI 05 analyzes the spin-independent  $X^0$ -nucleon cross section limits with both isoscalar and isovector couplings. See their Fig. 3 and 4 for limits on the couplings.

### For $m_{X^0} = 1 \text{ TeV}$

For limits from  $X^0$  annihilation in the Sun, the assumed annihilation final state is shown in parenthesis in the comment.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 8.6 \times 10^{-8}$	90	AGNES 16 DS50	Ar	
$< 2 \times 10^{-7}$	90	AGNES 15 DS1D	Ar	
$< 2 \times 10^{-7}$	90	<sup>1</sup> AGNESE 15b CDM2	Ge	
$< 1 \times 10^{-8}$	90	AKERIB 14 LUX	Xe	
$< 2.2 \times 10^{-6}$	90	<sup>2</sup> AVORIN 14 BAIK	H, solar $\nu$ ( $W^+ W^-$ )	
$< 5.5 \times 10^{-5}$	90	<sup>2</sup> AVORIN 14 BAIK	H, solar $\nu$ ( $b\bar{b}$ )	
$< 6.8 \times 10^{-7}$	90	<sup>2</sup> AVORIN 14 BAIK	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 3.46 \times 10^{-7}$	90	<sup>3</sup> AARTSEN 13 ICCB	H, solar $\nu$ ( $W^+ W^-$ )	
$< 7.75 \times 10^{-6}$	90	<sup>3</sup> AARTSEN 13 ICCB	H, solar $\nu$ ( $b\bar{b}$ )	
$< 6.9 \times 10^{-7}$	90	<sup>4</sup> ADRIAN-MAR.13 ANTR	H, solar $\nu$ ( $W^+ W^-$ )	
$< 1.5 \times 10^{-5}$	90	<sup>4</sup> ADRIAN-MAR.13 ANTR	H, solar $\nu$ ( $b\bar{b}$ )	
$< 1.8 \times 10^{-7}$	90	<sup>4</sup> ADRIAN-MAR.13 ANTR	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 4.3 \times 10^{-6}$	90	<sup>5</sup> BOLIEV 13 BAKS	H, solar $\nu$ ( $W^+ W^-$ )	
$< 3.4 \times 10^{-5}$	90	<sup>5</sup> BOLIEV 13 BAKS	H, solar $\nu$ ( $b\bar{b}$ )	
$< 1.2 \times 10^{-6}$	90	<sup>5</sup> BOLIEV 13 BAKS	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 2.12 \times 10^{-7}$	90	<sup>6</sup> ABBASI 12 ICCB	H, solar $\nu$ ( $W^+ W^-$ )	
$< 6.56 \times 10^{-6}$	90	<sup>6</sup> ABBASI 12 ICCB	H, solar $\nu$ ( $b\bar{b}$ )	
$< 4 \times 10^{-7}$	90	AKIMOV 12 ZEP3	Xe	
$< 1.1 \times 10^{-5}$	90	<sup>7</sup> ANGLORHER 12 CRES	CaWO <sub>4</sub>	
$< 2 \times 10^{-8}$	90	<sup>8</sup> APRILE 12 X100	Xe	
$< 2 \times 10^{-6}$	90	BEHNKE 12 COUP	CF <sub>3</sub> I	
$< 4 \times 10^{-6}$		FELIZARDO 12 SMPL	C <sub>2</sub> ClF <sub>5</sub>	
$< 1.5 \times 10^{-6}$	90	KIM 12 KIMS	Csl	
		<sup>9</sup> AHMED 11 CDM2	Ge, inelastic	
$< 1.5 \times 10^{-7}$	90	<sup>10</sup> AHMED 11A RVUE	Ge	
$< 2 \times 10^{-7}$	90	<sup>11</sup> APRILE 11 X100	Xe	
$< 8 \times 10^{-8}$	90	<sup>8</sup> APRILE 11b X100	Xe	
$< 2 \times 10^{-7}$	90	<sup>12</sup> ARMENGAUD 11 EDE2	Ge	
		<sup>13</sup> HORN 11 ZEP3	Xe	
$< 2 \times 10^{-7}$	90	AHMED 10 CDM2	Ge	
$< 4 \times 10^{-7}$	90	APRILE 10 X100	Xe	
$< 6 \times 10^{-7}$	90	ARMENGAUD 10 EDE2	Ge	
$< 3.5 \times 10^{-7}$	90	<sup>14</sup> AHMED 09 CDM2	Ge	

- <sup>1</sup> AGNESE 15b reanalyse AHMED 10 data.
- <sup>2</sup> AVORIN 14 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun in data taken between 1998 and 2003. See their Table 1 for limits assuming annihilation into neutrino pairs.
- <sup>3</sup> AARTSEN 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between June 2010 and May 2011.
- <sup>4</sup> ADRIAN-MARTINEZ 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between Jan. 2007 and Dec. 2008.
- <sup>5</sup> BOLIEV 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken from 1978 to 2009. See also SUVOROVA 13 for an older analysis of the same data.
- <sup>6</sup> ABBASI 12 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.
- <sup>7</sup> Reanalysis of ANGLORHER 09 data with all three nuclides. See also BROWN 12.
- <sup>8</sup> See also APRILE 14A.
- <sup>9</sup> AHMED 11 search for  $X^0$  inelastic scattering. See their Fig. 8–10 for limits.
- <sup>10</sup> AHMED 11A combine CDMs and EDELWEISS data.
- <sup>11</sup> APRILE 11 reanalyze APRILE 10 data.
- <sup>12</sup> Supersedes ARMENGAUD 10. A limit on inelastic cross section is also given.
- <sup>13</sup> HORN 11 perform detector calibration by neutrons. Earlier results are only marginally affected.
- <sup>14</sup> Superseded by AHMED 10.

### Limits for Spin-Dependent Cross Section of Dark Matter Particle ( $X^0$ ) on Proton

### For $m_{X^0} = 20 \text{ GeV}$

For limits from  $X^0$  annihilation in the Sun, the assumed annihilation final state is shown in parenthesis in the comment.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.2 \times 10^{-3}$	90	AMOLE 15 PICO	C <sub>3</sub> F <sub>8</sub>	
$< 1.43 \times 10^{-3}$	90	CHOI 15 SKAM	H, solar $\nu$ ( $b\bar{b}$ )	
$< 1.42 \times 10^{-4}$	90	CHOI 15 SKAM	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 5 \times 10^{-3}$	90	FELIZARDO 14 SMPL	C <sub>2</sub> ClF <sub>5</sub>	
$< 1.29 \times 10^{-2}$	90	<sup>1</sup> AARTSEN 13 ICCB	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 3.17 \times 10^{-2}$	90	<sup>2</sup> APRILE 13 X100	Xe	
$< 3 \times 10^{-2}$	90	ARCHAMBAU.12 PICA	F (C <sub>4</sub> F <sub>10</sub> )	

$< 6 \times 10^{-2}$	90	BEHNKE 12 COUP	CF <sub>3</sub> I	
$< 20$	90	DAW 12 DRFT	F (CF <sub>4</sub> )	
$< 7 \times 10^{-3}$		FELIZARDO 12 SMPL	C <sub>2</sub> ClF <sub>5</sub>	
$< 0.15$	90	KIM 12 KIMS	Csl	
$< 1 \times 10^5$	90	<sup>3</sup> AHLEN 11 DMTP	F (CF <sub>4</sub> )	
$< 0.1$	90	<sup>3</sup> BEHNKE 11 COUP	CF <sub>3</sub> I	
$< 1.5 \times 10^{-2}$	90	<sup>4</sup> TANAKA 11 SKAM	H, solar $\nu$ ( $b\bar{b}$ )	
$< 0.2$	90	ARCHAMBAU.09 PICA	F	
$< 4$	90	LEBEDENKO 09A ZEP3	Xe	
$< 0.6$	90	ANGLE 08A XE10	Xe	
$< 100$	90	ALNER 07 ZEP2	Xe	
$< 1$	90	LEE 07A KIMS	Csl	
$< 20$	90	<sup>5</sup> AKERIB 06 CDMs	<sup>73</sup> Ge, <sup>29</sup> Si	
$< 2$	90	SHIMIZU 06A CNTR	F (CaF <sub>2</sub> )	
$< 0.5$	90	ALNER 05 NAIA	NaI	
$< 1.5$	90	BARNABE-HE.05 PICA	F (C <sub>4</sub> F <sub>10</sub> )	
$< 1.5$	90	GIRARD 05 SMPL	F (C <sub>2</sub> ClF <sub>5</sub> )	
$< 35$	90	MUCHI 03 BOLO	LiF	
$< 30$	90	TAKEDA 03 BOLO	NaF	

- <sup>1</sup> AARTSEN 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between June 2010 and May 2011.
- <sup>2</sup> The value has been provided by the authors. APRILE 13 note that the proton limits on Xe are highly sensitive to the theoretical model used. See also APRILE 14A.
- <sup>3</sup> Use a direction-sensitive detector.
- <sup>4</sup> TANAKA 11 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.
- <sup>5</sup> See also AKERIB 05.

### For $m_{X^0} = 100 \text{ GeV}$

For limits from  $X^0$  annihilation in the Sun, the assumed annihilation final state is shown in parenthesis in the comment.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.5 \times 10^{-3}$	90	AMOLE 15 PICO	C <sub>3</sub> F <sub>8</sub>	
$< 3.19 \times 10^{-3}$	90	CHOI 15 SKAM	H, solar $\nu$ ( $b\bar{b}$ )	
$< 2.80 \times 10^{-4}$	90	CHOI 15 SKAM	H, solar $\nu$ ( $W^+ W^-$ )	
$< 1.24 \times 10^{-4}$	90	CHOI 15 SKAM	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 8 \times 10^2$	90	<sup>1</sup> NAKAMURA 15 NAGE	CF <sub>4</sub>	
$< 1.7 \times 10^{-3}$	90	<sup>2</sup> AVORIN 14 BAIK	H, solar $\nu$ ( $W^+ W^-$ )	
$< 4.5 \times 10^{-2}$	90	<sup>2</sup> AVORIN 14 BAIK	H, solar $\nu$ ( $b\bar{b}$ )	
$< 7.1 \times 10^{-4}$	90	<sup>2</sup> AVORIN 14 BAIK	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 6 \times 10^{-3}$	90	FELIZARDO 14 SMPL	C <sub>2</sub> ClF <sub>5</sub>	
$< 2.68 \times 10^{-4}$	90	<sup>3</sup> AARTSEN 13 ICCB	H, solar $\nu$ ( $W^+ W^-$ )	
$< 1.47 \times 10^{-2}$	90	<sup>3</sup> AARTSEN 13 ICCB	H, solar $\nu$ ( $b\bar{b}$ )	
$< 8.5 \times 10^{-4}$	90	<sup>4</sup> ADRIAN-MAR.13 ANTR	H, solar $\nu$ ( $W^+ W^-$ )	
$< 5.5 \times 10^{-2}$	90	<sup>4</sup> ADRIAN-MAR.13 ANTR	H, solar $\nu$ ( $b\bar{b}$ )	
$< 3.4 \times 10^{-4}$	90	<sup>4</sup> ADRIAN-MAR.13 ANTR	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 1.00 \times 10^{-2}$	90	<sup>5</sup> APRILE 13 X100	Xe	
$< 7.1 \times 10^{-4}$	90	<sup>6</sup> BOLIEV 13 BAKS	H, solar $\nu$ ( $W^+ W^-$ )	
$< 8.4 \times 10^{-3}$	90	<sup>6</sup> BOLIEV 13 BAKS	H, solar $\nu$ ( $b\bar{b}$ )	
$< 3.1 \times 10^{-4}$	90	<sup>6</sup> BOLIEV 13 BAKS	H, solar $\nu$ ( $\tau^+ \tau^-$ )	
$< 7.07 \times 10^{-4}$	90	<sup>7</sup> ABBASI 12 ICCB	H, solar $\nu$ ( $W^+ W^-$ )	
$< 4.53 \times 10^{-2}$	90	<sup>7</sup> ABBASI 12 ICCB	H, solar $\nu$ ( $b\bar{b}$ )	
$< 7 \times 10^{-2}$	90	ARCHAMBAU.12 PICA	F (C <sub>4</sub> F <sub>10</sub> )	
$< 1 \times 10^{-2}$	90	BEHNKE 12 COUP	CF <sub>3</sub> I	
$< 1.8$	90	DAW 12 DRFT	F (CF <sub>4</sub> )	
$< 9 \times 10^{-3}$		FELIZARDO 12 SMPL	C <sub>2</sub> ClF <sub>5</sub>	
$< 2 \times 10^{-2}$	90	KIM 12 KIMS	Csl	
$< 2 \times 10^3$	90	<sup>1</sup> AHLEN 11 DMTP	F (CF <sub>4</sub> )	
$< 7 \times 10^{-2}$	90	BEHNKE 11 COUP	CF <sub>3</sub> I	
$< 2.7 \times 10^{-4}$	90	<sup>8</sup> TANAKA 11 SKAM	H, solar $\nu$ ( $W^+ W^-$ )	
$< 4.5 \times 10^{-3}$	90	<sup>8</sup> TANAKA 11 SKAM	H, solar $\nu$ ( $b\bar{b}$ )	
		<sup>9</sup> FELIZARDO 10 SMPL	C <sub>2</sub> ClF <sub>3</sub>	
$< 6 \times 10^3$	90	<sup>1</sup> MUCHI 10 NAGE	CF <sub>4</sub>	
$< 0.4$	90	ARCHAMBAU.09 PICA	F	
$< 0.8$	90	LEBEDENKO 09A ZEP3	Xe	
$< 1.0$	90	ANGLE 08A XE10	Xe	
$< 15$	90	ALNER 07 ZEP2	Xe	
$< 0.2$	90	LEE 07A KIMS	Csl	
$< 1 \times 10^4$	90	<sup>1</sup> MUCHI 07 NAGE	F (CF <sub>4</sub> )	
$< 5$	90	<sup>10</sup> AKERIB 06 CDMs	<sup>73</sup> Ge, <sup>29</sup> Si	
$< 2$	90	SHIMIZU 06A CNTR	F (CaF <sub>2</sub> )	
$< 0.3$	90	ALNER 05 NAIA	NaI	
$< 2$	90	BARNABE-HE.05 PICA	F (C <sub>4</sub> F <sub>10</sub> )	
$< 100$	90	BENOIT 05 EDEL	<sup>73</sup> Ge	
$< 1.5$	90	GIRARD 05 SMPL	F (C <sub>2</sub> ClF <sub>5</sub> )	
$< 0.7$		<sup>11</sup> GIULIANI 05A RVUE		
		<sup>12</sup> GIULIANI 04 RVUE		
		<sup>13</sup> GIULIANI 04A RVUE		
$< 35$	90	MUCHI 03 BOLO	LiF	
$< 40$	90	TAKEDA 03 BOLO	NaF	

See key on page 601

# Searches Particle Listings

## WIMPs and Other Particle Searches

<sup>1</sup> Use a direction-sensitive detector.

<sup>2</sup> AVORIN 14 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun in data taken between 1998 and 2003. See their Table 1 for limits assuming annihilation into neutrino pairs.

<sup>3</sup> AARTSEN 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between June 2010 and May 2011.

<sup>4</sup> ADRIAN-MARTINEZ 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between Jan. 2007 and Dec. 2008.

<sup>5</sup> The value has been provided by the authors. APRILE 13 note that the proton limits on Xe are highly sensitive to the theoretical model used. See also APRILE 14A.

<sup>6</sup> BOLIEV 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken from 1978 to 2009. See also SUVOROVA 13 for an older analysis of the same data.

<sup>7</sup> ABBASI 12 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.

<sup>8</sup> TANAKA 11 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.

<sup>9</sup> See their Fig. 3 for limits on spin-dependent proton couplings for  $X^0$  mass of 50 GeV.

<sup>10</sup> See also AKERIB 05.

<sup>11</sup> GIULIANI 05A analyze available data and give combined limits.

<sup>12</sup> GIULIANI 04 reanalyze COLLAR 00 data and give limits for spin-dependent  $X^0$ -proton coupling.

<sup>13</sup> GIULIANI 04A give limits for spin-dependent  $X^0$ -proton couplings from existing data.

### For $m_{X^0} = 1 \text{ TeV}$

For limits from  $X^0$  annihilation in the Sun, the assumed annihilation final state is shown in parenthesis in the comment.

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< $1 \times 10^{-2}$	90	AMOLE 15	PICO	$C_3F_8$
< $1.5 \times 10^3$	90	NAKAMURA 15	NAGE	$CF_4$
< $2.7 \times 10^{-3}$	90	<sup>1</sup> AVORIN 14	BAIK	H, solar $\nu$ ( $W^+ W^-$ )
< $6.9 \times 10^{-2}$	90	<sup>1</sup> AVORIN 14	BAIK	H, solar $\nu$ ( $b\bar{b}$ )
< $8.4 \times 10^{-4}$	90	<sup>1</sup> AVORIN 14	BAIK	H, solar $\nu$ ( $\tau^+ \tau^-$ )
< $4.48 \times 10^{-4}$	90	<sup>2</sup> AARTSEN 13	ICCB	H, solar $\nu$ ( $W^+ W^-$ )
< $1.00 \times 10^{-2}$	90	<sup>2</sup> AARTSEN 13	ICCB	H, solar $\nu$ ( $b\bar{b}$ )
< $8.9 \times 10^{-4}$	90	<sup>3</sup> ADRIAN-MAR.13	ANTR	H, solar $\nu$ ( $W^+ W^-$ )
< $2.0 \times 10^{-2}$	90	<sup>3</sup> ADRIAN-MAR.13	ANTR	H, solar $\nu$ ( $b\bar{b}$ )
< $2.3 \times 10^{-4}$	90	<sup>3</sup> ADRIAN-MAR.13	ANTR	H, solar $\nu$ ( $\tau^+ \tau^-$ )
< $7.57 \times 10^{-2}$	90	<sup>4</sup> APRILE 13	X100	Xe
< $5.4 \times 10^{-3}$	90	<sup>5</sup> BOLIEV 13	BAKS	H, solar $\nu$ ( $W^+ W^-$ )
< $4.2 \times 10^{-2}$	90	<sup>5</sup> BOLIEV 13	BAKS	H, solar $\nu$ ( $b\bar{b}$ )
< $1.5 \times 10^{-3}$	90	<sup>5</sup> BOLIEV 13	BAKS	H, solar $\nu$ ( $\tau^+ \tau^-$ )
< $2.50 \times 10^{-4}$	90	<sup>6</sup> ABBASI 12	ICCB	H, solar $\nu$ ( $W^+ W^-$ )
< $7.86 \times 10^{-3}$	90	<sup>6</sup> ABBASI 12	ICCB	H, solar $\nu$ ( $b\bar{b}$ )
< $8 \times 10^{-2}$	90	BEHNKE 12	COUP	$CF_3I$
< 8	90	DAW 12	DRFT	F ( $CF_4$ )
< $6 \times 10^{-2}$	90	FELIZARDO 12	SMPPL	$C_2ClF_5$
< $8 \times 10^{-2}$	90	KIM 12	KIMS	Csl
< $8 \times 10^3$	90	<sup>7</sup> AHLEN 11	DMTP	F ( $CF_4$ )
< 0.4	90	BEHNKE 11	COUP	$CF_3I$
< $2 \times 10^{-3}$	90	<sup>8</sup> TANAKA 11	SKAM	H, solar $\nu$ ( $b\bar{b}$ )
< $2 \times 10^{-2}$	90	<sup>8</sup> TANAKA 11	SKAM	H, solar $\nu$ ( $W^+ W^-$ )
< $1 \times 10^{-3}$	90	<sup>9</sup> ABBASI 10	ICCB	KK dark matter
< $2 \times 10^4$	90	<sup>7</sup> MIUCHI 10	NAGE	$CF_4$
< $8.7 \times 10^{-4}$	90	ABBASI 09B	ICCB	H, solar $\nu$ ( $W^+ W^-$ )
< $2.2 \times 10^{-2}$	90	ABBASI 09B	ICCB	H, solar $\nu$ ( $b\bar{b}$ )
< 3	90	ARCHAMBAUD.09	PICA	F
< 6	90	LEBEDENKO 09A	ZEP3	Xe
< 9	90	ANGLE 08A	XE10	Xe
<100	90	ALNER 07	ZEP2	Xe
< 0.8	90	LEE 07A	KIMS	Csl
< $4 \times 10^4$	90	<sup>7</sup> MIUCHI 07	NAGE	F ( $CF_4$ )
< 30	90	<sup>10</sup> AKERIB 06	CDMS	$^{73}Ge$ , $^{29}Si$
< 1.5	90	ALNER 05	NAIA	NaI
< 15	90	BARNABE-HE.05	PICA	F ( $C_4F_{10}$ )
<600	90	BENOIT 05	EDEL	$^{73}Ge$
< 10	90	GIRARD 05	SMPPL	F ( $C_2ClF_5$ )
<260	90	MIUCHI 03	BOLO	LiF
<150	90	TAKEDA 03	BOLO	NaF

<sup>1</sup> AVORIN 14 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun in data taken between 1998 and 2003. See their Table 1 for limits assuming annihilation into neutrino pairs.

<sup>2</sup> AARTSEN 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between June 2010 and May 2011.

<sup>3</sup> ADRIAN-MARTINEZ 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken between Jan. 2007 and Dec. 2008.

<sup>4</sup> The value has been provided by the authors. APRILE 13 note that the proton limits on Xe are highly sensitive to the theoretical model used. See also APRILE 14A.

<sup>5</sup> BOLIEV 13 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the sun in data taken from 1978 to 2009. See also SUVOROVA 13 for an older analysis of the same data.

<sup>6</sup> ABBASI 12 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.

<sup>7</sup> Use a direction-sensitive detector.

<sup>8</sup> TANAKA 11 search for neutrinos from the Sun arising from the pair annihilation of  $X^0$  trapped by the Sun. The amount of  $X^0$  depends on the  $X^0$ -proton cross section.

<sup>9</sup> ABBASI 10 search for  $\nu_\mu$  from annihilations of Kaluza-Klein photon dark matter in the Sun.

<sup>10</sup> See also AKERIB 05.

### Limits for Spin-Dependent Cross Section of Dark Matter Particle ( $X^0$ ) on Neutron

#### For $m_{X^0} = 20 \text{ GeV}$

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.09	90	FELIZARDO 14	SMPPL	$C_2ClF_5$
< 8	90	<sup>1</sup> UCHIDA 14	XMAS	$^{129}Xe$ , inelastic
< $1.13 \times 10^{-3}$	90	<sup>2</sup> APRILE 13	X100	Xe
< 0.02	90	AKIMOV 12	ZEP3	Xe
		<sup>3</sup> AHMED 11B	CDM2	Ge, low threshold
< 0.06	90	AHMED 09	CDM2	Ge
< 0.04	90	LEBEDENKO 09A	ZEP3	Xe
< 50		<sup>4</sup> LIN 09	TEXO	Ge
< $6 \times 10^{-3}$	90	ANGLE 08A	XE10	Xe
< 0.5	90	ALNER 07	ZEP2	Xe
< 25	90	LEE 07A	KIMS	Csl
< 0.3	90	<sup>5</sup> AKERIB 06	CDMS	$^{73}Ge$ , $^{29}Si$
< 30	90	SHIMIZU 06A	CNTR	F ( $CaF_2$ )
< 60	90	ALNER 05	NAIA	NaI
< 20	90	BARNABE-HE.05	PICA	F ( $C_4F_{10}$ )
< 10	90	BENOIT 05	EDEL	$^{73}Ge$
< 4	90	KLAPDOR-K...05	HDMS	$^{73}Ge$ (enriched)
<600	90	TAKEDA 03	BOLO	NaF

<sup>1</sup> Derived limit from search for inelastic scattering  $X^0 + ^{129}Xe \rightarrow X^0 + ^{129}Xe^*$  (39.58 keV).

<sup>2</sup> The value has been provided by the authors. See also APRILE 14A.

<sup>3</sup> AHMED 11B give limits on spin-dependent  $X^0$ -neutron cross section for  $m_{X^0} = 4-12 \text{ GeV}$  in the range  $10^{-3}-10 \text{ pb}$ . See their Fig. 3.

<sup>4</sup> See their Fig. 6(b) for cross section limits for  $m_{X^0}$  extending down to 2 GeV.

<sup>5</sup> See also AKERIB 05.

#### For $m_{X^0} = 100 \text{ GeV}$

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.1	90	FELIZARDO 14	SMPPL	$C_2ClF_5$
< 0.05	90	<sup>1</sup> UCHIDA 14	XMAS	$^{129}Xe$ , inelastic
< $4.68 \times 10^{-4}$	90	<sup>2</sup> APRILE 13	X100	Xe
< 0.01	90	AKIMOV 12	ZEP3	Xe
		<sup>3</sup> FELIZARDO 10	SMPPL	$C_2ClF_3$
< 0.02	90	AHMED 09	CDM2	Ge
< 0.01	90	LEBEDENKO 09A	ZEP3	Xe
<100		LIN 09	TEXO	Ge
< 0.01	90	ANGLE 08A	XE10	Xe
< 0.05	90	<sup>4</sup> BEDNYAKOV 08	RVUE	Ge
< 0.08	90	ALNER 07	ZEP2	Xe
< 6	90	LEE 07A	KIMS	Csl
< 0.07	90	<sup>5</sup> AKERIB 06	CDMS	$^{73}Ge$ , $^{29}Si$
< 30	90	SHIMIZU 06A	CNTR	F ( $CaF_2$ )
< 10	90	ALNER 05	NAIA	NaI
< 30	90	BARNABE-HE.05	PICA	F ( $C_4F_{10}$ )
< 0.7	90	BENOIT 05	EDEL	$^{73}Ge$
< 0.2		<sup>6</sup> GIULIANI 05A	RVUE	
< 1.5	90	KLAPDOR-K...05	HDMS	$^{73}Ge$ (enriched)
		<sup>7</sup> GIULIANI 04	RVUE	
		<sup>8</sup> GIULIANI 04A	RVUE	
		<sup>9</sup> MIUCHI 03	BOLO	LiF
<800	90	TAKEDA 03	BOLO	NaF

<sup>1</sup> Derived limit from search for inelastic scattering  $X^0 + ^{129}Xe \rightarrow X^0 + ^{129}Xe^*$  (39.58 keV).

<sup>2</sup> The value has been provided by the authors. See also APRILE 14A.

<sup>3</sup> See their Fig. 3 for limits on spin-dependent neutron couplings for  $X^0$  mass of 50 GeV.

<sup>4</sup> BEDNYAKOV 08 reanalyze KLAPDOR-KLEINGROTHAUS 05 and BAUDIS 01 data.

<sup>5</sup> See also AKERIB 05.

<sup>6</sup> GIULIANI 05A analyze available data and give combined limits.

<sup>7</sup> GIULIANI 04 reanalyze COLLAR 00 data and give limits for spin-dependent  $X^0$ -neutron coupling.

<sup>8</sup> GIULIANI 04A give limits for spin-dependent  $X^0$ -neutron couplings from existing data.

<sup>9</sup> MIUCHI 03 give model-independent limit for spin-dependent  $X^0$ -proton and neutron cross sections. See their Fig. 5.

#### For $m_{X^0} = 1 \text{ TeV}$

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.07	90	FELIZARDO 14	SMPPL	$C_2ClF_5$
< 0.2	90	<sup>1</sup> UCHIDA 14	XMAS	$^{129}Xe$ , inelastic
< $3.64 \times 10^{-3}$	90	<sup>2</sup> APRILE 13	X100	Xe
< 0.08	90	AKIMOV 12	ZEP3	Xe
< 0.2	90	AHMED 09	CDM2	Ge
< 0.1	90	LEBEDENKO 09A	ZEP3	Xe
< 0.1	90	ANGLE 08A	XE10	Xe
< 0.25	90	<sup>3</sup> BEDNYAKOV 08	RVUE	Ge

# Searches Particle Listings

## WIMPs and Other Particle Searches

< 0.6	90	ALNER	07	ZEP2	Xe
< 30	90	LEE	07A	KIMS	Csl
< 0.5	90	4 AKERIB	06	CDMS	<sup>73</sup> Ge, <sup>29</sup> Si
< 40	90	ALNER	05	NAIA	NaI
<200	90	BARNABE-HE...	05	PICA	F (C <sub>4</sub> F <sub>10</sub> )
< 4	90	BENOIT	05	EDEL	<sup>73</sup> Ge
< 10	90	KLAPDOR-K...	05	HDMS	<sup>73</sup> Ge (enriched)
< 4 × 10 <sup>3</sup>	90	TAKEDA	03	BOLO	NaF

<sup>1</sup> Derived limit from search for inelastic scattering  $\chi^0 + ^{129}\text{Xe}^* \rightarrow \chi^0 + ^{129}\text{Xe}^*$  (39.58 keV).

<sup>2</sup> The value has been provided by the authors. See also APRILE 14A.

<sup>3</sup> BEDNYAKOV 08 reanalyzed KLAPDOR-KLEINGROTHAUS 05 and BAUDIS 01 data.

<sup>4</sup> See also AKERIB 05.

### Cross-Section Limits for Dark Matter Particles ( $\chi^0$ ) on Nuclei

For  $m_{\chi^0} = 20 \text{ GeV}$

VALUE (nb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.03	90	1 UCHIDA	14	XMAS <sup>129</sup> Xe, inelastic
< 0.08	90	2 ANGLOHER	02	CRES Al
		3 BENOIT	00	EDEL Ge
< 0.04	95	4 KLIMENKO	98	CNTR <sup>73</sup> Ge, inel.
< 0.8		ALESSAND...	96	CNTR O
< 6		ALESSAND...	96	CNTR Te
< 0.02	90	5 BELLI	96	CNTR <sup>129</sup> Xe, inel.
		6 BELLI	96c	CNTR <sup>129</sup> Xe
< 4 × 10 <sup>-3</sup>	90	7 BERNABEI	96	CNTR Na
< 0.3	90	7 BERNABEI	96	CNTR I
< 0.2	95	8 SARSA	96	CNTR Na
< 0.015	90	9 SMITH	96	CNTR Na
< 0.05	95	10 GARCIA	95	CNTR Natural Ge
< 0.1	95	QUENBY	95	CNTR Na
<90	90	11 SNOWDEN...	95	MICA <sup>16</sup> O
< 4 × 10 <sup>3</sup>	90	11 SNOWDEN...	95	MICA <sup>39</sup> K
< 0.7	90	BACCI	92	CNTR Na
< 0.12	90	12 REUSSER	91	CNTR Natural Ge
< 0.06	95	CALDWELL	88	CNTR Natural Ge

<sup>1</sup> UCHIDA 14 limit is for inelastic scattering  $\chi^0 + ^{129}\text{Xe}^* \rightarrow \chi^0 + ^{129}\text{Xe}^*$  (39.58 keV).

<sup>2</sup> ANGLOHER 02 limit is for spin-dependent WIMP-Aluminum cross section.

<sup>3</sup> BENOIT 00 find four event categories in Ge detectors and suggest that low-energy surface nuclear recoils can explain anomalous events reported by UKDMC and Saclay NaI experiments.

<sup>4</sup> KLIMENKO 98 limit is for inelastic scattering  $\chi^0 \text{ } ^{73}\text{Ge} \rightarrow \chi^0 \text{ } ^{73}\text{Ge}^*$  (13.26 keV).

<sup>5</sup> BELLI 96 limit for inelastic scattering  $\chi^0 \text{ } ^{129}\text{Xe} \rightarrow \chi^0 \text{ } ^{129}\text{Xe}^*$  (39.58 keV).

<sup>6</sup> BELLI 96c use background subtraction and obtain  $\sigma < 150 \text{ pb}$  ( $< 1.5 \text{ fb}$ ) (90% CL) for spin-dependent (independent)  $\chi^0$ -proton cross section. The confidence level is from R. Bernabei, private communication, May 20, 1999.

<sup>7</sup> BERNABEI 96 use pulse shape discrimination to enhance the possible signal. The limit here is from R. Bernabei, private communication, September 19, 1997.

<sup>8</sup> SARSA 96 search for annual modulation of WIMP signal. See SARSA 97 for details of the analysis. The limit here is from M.L. Sarsa, private communication, May 26, 1997.

<sup>9</sup> SMITH 96 use pulse shape discrimination to enhance the possible signal. A dark matter density of  $0.4 \text{ GeV cm}^{-3}$  is assumed.

<sup>10</sup> GARCIA 95 limit is from the event rate. A weaker limit is obtained from searches for diurnal and annual modulation.

<sup>11</sup> SNOWDEN-IFFT 95 look for recoil tracks in an ancient mica crystal. Similar limits are also given for <sup>27</sup>Al and <sup>28</sup>Si. See COLLAR 96 and SNOWDEN-IFFT 96 for discussion on potential backgrounds.

<sup>12</sup> REUSSER 91 limit here is changed from published (0.04) after reanalysis by authors. J.L. Vuilleumier, private communication, March 29, 1996.

For  $m_{\chi^0} = 100 \text{ GeV}$

VALUE (nb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 3 × 10 <sup>-3</sup>	90	1 UCHIDA	14	XMAS <sup>129</sup> Xe, inelastic
< 0.3	90	2 ANGLOHER	02	CRES Al
		3 BELLI	02	RVUE
		4 BERNABEI	02c	DAMA
		5 GREEN	02	RVUE
		6 ULLIO	01	RVUE
		7 BENOIT	00	EDEL Ge
< 4 × 10 <sup>-3</sup>	90	8 BERNABEI	00d	<sup>129</sup> Xe, inel.
		9 AMBROSIO	99	MCRO
		10 BRHLIK	99	RVUE
< 8 × 10 <sup>-3</sup>	95	11 KLIMENKO	98	CNTR <sup>73</sup> Ge, inel.
< 0.08	95	12 KLIMENKO	98	CNTR <sup>73</sup> Ge, inel.
< 4		ALESSAND...	96	CNTR O
<25		ALESSAND...	96	CNTR Te
< 6 × 10 <sup>-3</sup>	90	13 BELLI	96	CNTR <sup>129</sup> Xe, inel.
		14 BELLI	96c	CNTR <sup>129</sup> Xe
< 1 × 10 <sup>-3</sup>	90	15 BERNABEI	96	CNTR Na
< 0.3	90	15 BERNABEI	96	CNTR I

< 0.7	95	16 SARSA	96	CNTR Na
< 0.03	90	17 SMITH	96	CNTR Na
< 0.8	90	17 SMITH	96	CNTR I
< 0.35	95	18 GARCIA	95	CNTR Natural Ge
< 0.6	95	QUENBY	95	CNTR Na
< 3	95	QUENBY	95	CNTR I
< 1.5 × 10 <sup>2</sup>	90	19 SNOWDEN...	95	MICA <sup>16</sup> O
< 4 × 10 <sup>2</sup>	90	19 SNOWDEN...	95	MICA <sup>39</sup> K
< 0.08	90	20 BECK	94	CNTR <sup>76</sup> Ge
< 2.5	90	BACCI	92	CNTR Na
< 3	90	BACCI	92	CNTR I
< 0.9	90	21 REUSSER	91	CNTR Natural Ge
< 0.7	95	CALDWELL	88	CNTR Natural Ge

<sup>1</sup> UCHIDA 14 limit is for inelastic scattering  $\chi^0 + ^{129}\text{Xe}^* \rightarrow \chi^0 + ^{129}\text{Xe}^*$  (39.58 keV).

<sup>2</sup> ANGLOHER 02 limit is for spin-dependent WIMP-Aluminum cross section.

<sup>3</sup> BELLI 02 discuss dependence of the extracted WIMP cross section on the assumptions of the galactic halo structure.

<sup>4</sup> BERNABEI 02c analyze the DAMA data in the scenario in which  $\chi^0$  scatters into a slightly heavier state as discussed by SMITH 01.

<sup>5</sup> GREEN 02 discusses dependence of extracted WIMP cross section limits on the assumptions of the galactic halo structure.

<sup>6</sup> ULLIO 01 disfavor the possibility that the BERNABEI 99 signal is due to spin-dependent WIMP coupling.

<sup>7</sup> BENOIT 00 find four event categories in Ge detectors and suggest that low-energy surface nuclear recoils can explain anomalous events reported by UKDMC and Saclay NaI experiments.

<sup>8</sup> BERNABEI 00d limit is for inelastic scattering  $\chi^0 \text{ } ^{129}\text{Xe} \rightarrow \chi^0 \text{ } ^{129}\text{Xe}^*$  (39.58 keV).

<sup>9</sup> AMBROSIO 99 search for upgoing muon events induced by neutrinos originating from WIMP annihilations in the Sun and Earth.

<sup>10</sup> BRHLIK 99 discuss the effect of astrophysical uncertainties on the WIMP interpretation of the BERNABEI 99 signal.

<sup>11</sup> KLIMENKO 98 limit is for inelastic scattering  $\chi^0 \text{ } ^{73}\text{Ge} \rightarrow \chi^0 \text{ } ^{73}\text{Ge}^*$  (13.26 keV).

<sup>12</sup> KLIMENKO 98 limit is for inelastic scattering  $\chi^0 \text{ } ^{73}\text{Ge} \rightarrow \chi^0 \text{ } ^{73}\text{Ge}^*$  (66.73 keV).

<sup>13</sup> BELLI 96 limit for inelastic scattering  $\chi^0 \text{ } ^{129}\text{Xe} \rightarrow \chi^0 \text{ } ^{129}\text{Xe}^*$  (39.58 keV).

<sup>14</sup> BELLI 96c use background subtraction and obtain  $\sigma < 0.35 \text{ pb}$  ( $< 0.15 \text{ fb}$ ) (90% CL) for spin-dependent (independent)  $\chi^0$ -proton cross section. The confidence level is from R. Bernabei, private communication, May 20, 1999.

<sup>15</sup> BERNABEI 96 use pulse shape discrimination to enhance the possible signal. The limit here is from R. Bernabei, private communication, September 19, 1997.

<sup>16</sup> SARSA 96 search for annual modulation of WIMP signal. See SARSA 97 for details of the analysis. The limit here is from M.L. Sarsa, private communication, May 26, 1997.

<sup>17</sup> SMITH 96 use pulse shape discrimination to enhance the possible signal. A dark matter density of  $0.4 \text{ GeV cm}^{-3}$  is assumed.

<sup>18</sup> GARCIA 95 limit is from the event rate. A weaker limit is obtained from searches for diurnal and annual modulation.

<sup>19</sup> SNOWDEN-IFFT 95 look for recoil tracks in an ancient mica crystal. Similar limits are also given for <sup>27</sup>Al and <sup>28</sup>Si. See COLLAR 96 and SNOWDEN-IFFT 96 for discussion on potential backgrounds.

<sup>20</sup> BECK 94 uses enriched <sup>76</sup>Ge (86% purity).

<sup>21</sup> REUSSER 91 limit here is changed from published (0.3) after reanalysis by authors. J.L. Vuilleumier, private communication, March 29, 1996.

For  $m_{\chi^0} = 1 \text{ TeV}$

VALUE (nb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.03	90	1 UCHIDA	14	XMAS <sup>129</sup> Xe, inelastic
< 3	90	2 ANGLOHER	02	CRES Al
		3 BENOIT	00	EDEL Ge
		4 BERNABEI	99d	CNTR SIMP
		5 DERBIN	99	CNTR SIMP
< 0.06	95	6 KLIMENKO	98	CNTR <sup>73</sup> Ge, inel.
< 0.4	95	7 KLIMENKO	98	CNTR <sup>73</sup> Ge, inel.
< 40		ALESSAND...	96	CNTR O
<700		ALESSAND...	96	CNTR Te
< 0.05	90	8 BELLI	96	CNTR <sup>129</sup> Xe, inel.
< 1.5	90	9 BELLI	96	CNTR <sup>129</sup> Xe, inel.
		10 BELLI	96c	CNTR <sup>129</sup> Xe
< 0.01	90	11 BERNABEI	96	CNTR Na
< 9	90	11 BERNABEI	96	CNTR I
< 7	95	12 SARSA	96	CNTR Na
< 0.3	90	13 SMITH	96	CNTR Na
< 6	90	13 SMITH	96	CNTR I
< 6	95	14 GARCIA	95	CNTR Natural Ge
< 8	95	QUENBY	95	CNTR Na
< 50	95	QUENBY	95	CNTR I
<700	90	15 SNOWDEN...	95	MICA <sup>16</sup> O
< 1 × 10 <sup>3</sup>	90	15 SNOWDEN...	95	MICA <sup>39</sup> K
< 0.8	90	16 BECK	94	CNTR <sup>76</sup> Ge
< 30	90	BACCI	92	CNTR Na
< 30	90	BACCI	92	CNTR I
< 15	90	17 REUSSER	91	CNTR Natural Ge
< 6	95	CALDWELL	88	CNTR Natural Ge

See key on page 601

## Searches Particle Listings

### WIMPs and Other Particle Searches

- <sup>1</sup> UCHIDA 14 limit is for inelastic scattering  $X^0 + {}^{129}\text{Xe}^* \rightarrow X^0 + {}^{129}\text{Xe}^*$  (39.58 keV).
- <sup>2</sup> ANGLÖHER 02 limit is for spin-dependent WIMP-Aluminum cross section.
- <sup>3</sup> BENOIT 00 find four event categories in Ge detectors and suggest that low-energy surface nuclear recoils can explain anomalous events reported by UKDMC and Saclay Nal experiments.
- <sup>4</sup> BERNABEI 99d search for SIMPs (Strongly Interacting Massive Particles) in the mass range  $10^3\text{--}10^{16}$  GeV. See their Fig. 3 for cross-section limits.
- <sup>5</sup> DERBIN 99 search for SIMPs (Strongly Interacting Massive Particles) in the mass range  $10^2\text{--}10^{14}$  GeV. See their Fig. 3 for cross-section limits.
- <sup>6</sup> KLIMENKO 98 limit is for inelastic scattering  $X^0 {}^{73}\text{Ge} \rightarrow X^0 {}^{73}\text{Ge}^*$  (13.26 keV).
- <sup>7</sup> KLIMENKO 98 limit is for inelastic scattering  $X^0 {}^{73}\text{Ge} \rightarrow X^0 {}^{73}\text{Ge}^*$  (66.73 keV).
- <sup>8</sup> BELL 96 limit for inelastic scattering  $X^0 {}^{129}\text{Xe} \rightarrow X^0 {}^{129}\text{Xe}^*$  (39.58 keV).
- <sup>9</sup> BELL 96 limit for inelastic scattering  $X^0 {}^{129}\text{Xe} \rightarrow X^0 {}^{129}\text{Xe}^*$  (236.14 keV).
- <sup>10</sup> BELL 96c use background subtraction and obtain  $\sigma < 0.7$  pb ( $< 0.7$  fb) (90% CL) for spin-dependent (independent)  $X^0$ -proton cross section. The confidence level is from R. Bernabei, private communication, May 20, 1999.
- <sup>11</sup> BERNABEI 96 use pulse shape discrimination to enhance the possible signal. The limit here is from R. Bernabei, private communication, September 19, 1997.
- <sup>12</sup> SARSA 96 search for annual modulation of WIMP. See SARSA 97 for details of the analysis. The limit here is from M.L. Sarsa, private communication, May 26, 1997.
- <sup>13</sup> SMITH 96 use pulse shape discrimination to enhance the possible signal. A dark matter density of  $0.4 \text{ GeV cm}^{-3}$  is assumed.
- <sup>14</sup> GARCIA 95 limit is from the event rate. A weaker limit is obtained from searches for diurnal and annual modulation.
- <sup>15</sup> SNOWDEN-IFFT 95 look for recoil tracks in an ancient mica crystal. Similar limits are also given for  ${}^{27}\text{Al}$  and  ${}^{28}\text{Si}$ . See COLLAR 96 and SNOWDEN-IFFT 96 for discussion on potential backgrounds.
- <sup>16</sup> BECK 94 uses enriched  ${}^{76}\text{Ge}$  (86% purity).
- <sup>17</sup> REUSSER 91 limit here is changed from published (5) after reanalysis by authors. J.L. Vuilleumier, private communication, March 29, 1996.

#### Miscellaneous Results from Underground Dark Matter Searches

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	1 APRILE	15	X100 Event rate modulation
	2 APRILE	15A	X100 Electron scattering
<sup>1</sup> APRILE 15 search for periodic variation of electronic recoil event rate in the data between Feb. 2011 and Mar. 2012. No significant modulation is found for periods up to 500 days.			
<sup>2</sup> APRILE 15A search for $X^0$ scattering off electrons. See their Fig. 4 for limits on cross section through axial-vector coupling for $m_{X^0}$ between 0.6 GeV and 1 TeV. For $m_{X^0} = 2$ GeV, $\sigma < 60$ pb (90%CL) is obtained.			

#### $X^0$ Annihilation Cross Section

Limits are on  $\sigma v$  for  $X^0$  pair annihilation at threshold.

VALUE ( $\text{cm}^3\text{s}^{-1}$ )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		1 AARTSEN	15c	ICCB $\nu$ , Galactic halo
		2 AARTSEN	15E	ICCB $\nu$ , Galactic center
		3 ABRAMOWSKI15	HESS	Galactic center
		4 ACKERMANN 15	FLAT	monochromatic $\gamma$
		5 ACKERMANN 15A	FLAT	isotropic $\gamma$ background
		6 ACKERMANN 15B	FLAT	Satellite galaxy
		7 ADRIAN-MAR.15	ANTR	$\nu$ , Galactic center
$< 2.90 \times 10^{-26}$	95	8,9 ACKERMANN 14	FLAT	Satellite galaxy, $m = 10$ GeV
$< 1.84 \times 10^{-25}$	95	8,10 ACKERMANN 14	FLAT	Satellite galaxy, $m = 100$ GeV
$< 1.75 \times 10^{-24}$	95	8,10 ACKERMANN 14	FLAT	Satellite galaxy, $m = 1$ TeV
$< 4.52 \times 10^{-24}$	95	11 ALEKSIC 14	MGIC	Segue 1, $m = 1.35$ TeV
		12 AARTSEN 13c	ICCB	Galaxies
		13 ABRA MOWSKI13	HESS	Central Galactic Halo
		14 ACKERMANN 13A	FLAT	Galaxy
		15 ABRA MOWSKI12	HESS	Fornax Cluster
		16 ACKERMANN 12	FLAT	Galaxy
		17 ACKERMANN 12	FLAT	Galaxy
		18 ALIU 12	VRTS	Segue 1
$< 1 \times 10^{-22}$	90	19 ABBASI 11c	ICCB	Galactic halo, $m=1$ TeV
$< 3 \times 10^{-25}$	95	20 ABRA MOWSKI11	HESS	Near Galactic center, $m=1$ TeV
$< 1 \times 10^{-26}$	95	21 ACKERMANN 11	FLAT	Satellite galaxy, $m=10$ GeV
$< 1 \times 10^{-25}$	95	21 ACKERMANN 11	FLAT	Satellite galaxy, $m=100$ GeV
$< 1 \times 10^{-24}$	95	21 ACKERMANN 11	FLAT	Satellite galaxy, $m=1$ TeV

- <sup>1</sup> AARTSEN 15c search for neutrinos from  $X^0$  annihilation in the Galactic halo. See their Figs. 16 and 17, and Table 5 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 100 GeV and 100 TeV.
- <sup>2</sup> AARTSEN 15E search for neutrinos from  $X^0$  annihilation in the Galactic center. See their Figs. 7 and 9, and Table 3 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 30 GeV and 10 TeV.
- <sup>3</sup> ABRAMOWSKI 15 search for  $\gamma$  from  $X^0$  annihilation in the Galactic center. See their Fig. 4 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 250 GeV and 10 TeV.
- <sup>4</sup> ACKERMANN 15 search for monochromatic  $\gamma$  from  $X^0$  annihilation in the Galactic halo. See their Fig. 8 and Tables 2-4 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 0.2 GeV and 500 GeV.
- <sup>5</sup> ACKERMANN 15A search for  $\gamma$  from  $X^0$  annihilation (both Galactic and extragalactic) in the isotropic  $\gamma$  background. See their Fig. 7 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 10 GeV and 30 TeV.
- <sup>6</sup> ACKERMANN 15B search for  $\gamma$  from  $X^0$  annihilation in 15 dwarf spheroidal satellite galaxies of the Milky Way. See their Figs. 1 and 2 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 2 GeV and 10 TeV.

- <sup>7</sup> ADRIAN-MARTINEZ 15 search for neutrinos from  $X^0$  annihilation in the Galactic center. See their Figs. 10 and 11 and Tables 1 and 2 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 25 GeV and 10 TeV.
- <sup>8</sup> ACKERMANN 14 search for  $\gamma$  from  $X^0$  annihilation in 25 dwarf spheroidal satellite galaxies of the Milky Way. See their Tables II-VII for limits assuming annihilation into  $e^+e^-$ ,  $\mu^+\mu^-$ ,  $\tau^+\tau^-$ ,  $u\bar{u}$ ,  $b\bar{b}$ , and  $W^+W^-$ , for  $X^0$  mass ranging from 2 GeV to 10 TeV.
- <sup>9</sup> Limit assuming  $X^0$  pair annihilation into  $b\bar{b}$ .
- <sup>10</sup> Limit assuming  $X^0$  pair annihilation into  $W^+W^-$ .
- <sup>11</sup> ALEKSIC 14 search for  $\gamma$  from  $X^0$  annihilation in the dwarf spheroidal galaxy Segue 1. The listed limit assumes annihilation into  $W^+W^-$ . See their Figs. 6, 7, and 16 for limits on  $\sigma \cdot v$  for annihilation channels  $\mu^+\mu^-$ ,  $\tau^+\tau^-$ ,  $b\bar{b}$ ,  $t\bar{t}$ ,  $\gamma\gamma$ ,  $\gamma Z$ ,  $W^+W^-$ ,  $ZZ$  for  $X^0$  mass between  $10^2$  and  $10^4$  GeV.
- <sup>12</sup> AARTSEN 13c search for neutrinos from  $X^0$  annihilation in nearby galaxies and galaxy clusters. See their Figs. 5-7 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 300 GeV and 100 TeV.
- <sup>13</sup> ABRAMOWSKI 13 search for monochromatic  $\gamma$  from  $X^0$  annihilation in the Milky Way halo in the central region. Limit on  $\sigma \cdot v$  between  $10^{-28}$  and  $10^{-25} \text{ cm}^3\text{s}^{-1}$  (95% CL) is obtained for  $X^0$  mass between 500 GeV and 20 TeV for  $X^0 \rightarrow \gamma\gamma$ .  $X^0$  density distribution in the Galaxy by Einasto is assumed. See their Fig. 4.
- <sup>14</sup> ACKERMANN 13A search for monochromatic  $\gamma$  from  $X^0$  annihilation in the Milky Way. Limit on  $\sigma \cdot v$  for the process  $X^0 X^0 \rightarrow \gamma\gamma$  in the range  $10^{-29}\text{--}10^{-27} \text{ cm}^3\text{s}^{-1}$  (95% CL) is obtained for  $X^0$  mass between 5 and 300 GeV. The limit depends slightly on the assumed density profile of  $X^0$  in the Galaxy. See their Tables VII-X and Fig.10. Supersedes ACKERMANN 12.
- <sup>15</sup> ABRAMOWSKI 12 search for  $\gamma$ 's from  $X^0$  annihilation in the Fornax galaxy cluster. See their Fig. 7 for limits on  $\sigma \cdot v$  for  $X^0$  mass between 0.1 and 100 TeV for the annihilation channels  $\tau^+\tau^-$ ,  $b\bar{b}$ , and  $W^+W^-$ .
- <sup>16</sup> ACKERMANN 12 search for monochromatic  $\gamma$  from  $X^0$  annihilation in the Milky Way. Limit on  $\sigma \cdot v$  in the range  $10^{-28}\text{--}10^{-26} \text{ cm}^3\text{s}^{-1}$  (95% CL) is obtained for  $X^0$  mass between 7 and 200 GeV if  $X^0$  annihilates into  $\gamma\gamma$ . The limit depends slightly on the assumed density profile of  $X^0$  in the Galaxy. See their Table III and Fig. 15.
- <sup>17</sup> ACKERMANN 12 search for  $\gamma$  from  $X^0$  annihilation in the Milky Way in the diffuse  $\gamma$  background. Limit on  $\sigma \cdot v$  of  $10^{-24} \text{ cm}^3\text{s}^{-1}$  or larger is obtained for  $X^0$  mass between 5 GeV and 10 TeV for various annihilation channels including  $W^+W^-$ ,  $b\bar{b}$ ,  $g g$ ,  $e^+e^-$ ,  $\mu^+\mu^-$ ,  $\tau^+\tau^-$ . The limit depends slightly on the assumed density profile of  $X^0$  in the Galaxy. See their Figs. 17-20.
- <sup>18</sup> ALIU 12 search for  $\gamma$ 's from  $X^0$  annihilation in the dwarf spheroidal galaxy Segue 1. Limit on  $\sigma \cdot v$  in the range  $10^{-24}\text{--}10^{-20} \text{ cm}^3\text{s}^{-1}$  (95% CL) is obtained for  $X^0$  mass between 10 GeV and 2 TeV for annihilation channels  $e^+e^-$ ,  $\mu^+\mu^-$ ,  $\tau^+\tau^-$ ,  $b\bar{b}$ , and  $W^+W^-$ . See their Fig. 3.
- <sup>19</sup> ABBASI 11c search for  $\nu_\mu$  from  $X^0$  annihilation in the outer halo of the Milky Way. The limit assumes annihilation into  $\nu\nu$ . See their Fig. 9 for limits with other annihilation channels.
- <sup>20</sup> ABRAMOWSKI 11 search for  $\gamma$  from  $X^0$  annihilation near the Galactic center. The limit assumes Einasto DM density profile.
- <sup>21</sup> ACKERMANN 11 search for  $\gamma$  from  $X^0$  annihilation in ten dwarf spheroidal satellite galaxies of the Milky Way. The limit for  $m = 10$  GeV assumes annihilation into  $b\bar{b}$ , the others  $W^+W^-$ . See their Fig. 2 for limits with other final states. See also GERINGER-SAMETH 11 for a different analysis of the same data.

#### Dark Matter Particle ( $X^0$ ) Production in Hadron Collisions

Searches for  $X^0$  production in association with observable particles ( $\gamma$ , jets, ...) in high energy hadron collisions. If a specific form of effective interaction Lagrangian is assumed, the limits may be translated into limits on  $X^0$ -nucleon scattering cross section.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	1 AAD	15As ATLS	$b(\bar{b}) + \cancel{E}_T$ , $t\bar{t} + \cancel{E}_T$
	2 AAD	15BH ATLS	jet + $\cancel{E}_T$
	3 AAD	15CF ATLS	$H^0 + \cancel{E}_T$
	4 AAD	15CS ATLS	$\gamma + \cancel{E}_T$
	5 KHACHATRY..15Ag CMS		$t\bar{t} + \cancel{E}_T$
	6 KHACHATRY..15AL CMS		jet + $\cancel{E}_T$
	7 KHACHATRY..15T CMS		$\ell + \cancel{E}_T$
	8 AAD	14Al ATLS	$W + \cancel{E}_T$
	9 AAD	14K ATLS	$Z + \cancel{E}_T$
	10 AAD	14O ATLS	$Z + \cancel{E}_T$
	11 AAD	13AD ATLS	jet + $\cancel{E}_T$
	12 AAD	13C ATLS	$\gamma + \cancel{E}_T$
	13 AALTONEN 12K CDF		$t + \cancel{E}_T$
	14 AALTONEN 12M CDF		jet + $\cancel{E}_T$
	15 CHATRCHYAN12AP CMS		jet + $\cancel{E}_T$
	16 CHATRCHYAN12T CMS		$\gamma + \cancel{E}_T$

- <sup>1</sup> AAD 15As search for events with one or more bottom quark and missing  $E_T$ , and also events with a top quark pair and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ . See their Figs. 5 and 6 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}700$  GeV.
- <sup>2</sup> AAD 15BH search for events with a jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ . See their Fig. 12 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}1200$  GeV.
- <sup>3</sup> AAD 15CF search for events with a  $H^0$  ( $\rightarrow \gamma\gamma$ ) and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ . See paper for limits on the strength of some contact interactions containing  $X^0$  and the Higgs fields.
- <sup>4</sup> AAD 15CS search for events with a photon and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ . See their Fig. 13 (see also erratum) for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}1000$  GeV.

# Searches Particle Listings

## WIMPs and Other Particle Searches

<sup>5</sup> KHACHATRYAN 15AG search for events with a top quark pair and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 19.7 \text{ fb}^{-1}$ . See their Fig. 8 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}200$  GeV.

<sup>6</sup> KHACHATRYAN 15AL search for events with a jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 19.7 \text{ fb}^{-1}$ . See their Fig. 5 and Tables 4–6 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}1000$  GeV.

<sup>7</sup> KHACHATRYAN 15T search for events with a lepton and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 19.7 \text{ fb}^{-1}$ . See their Fig. 17 for translated limits on  $X^0$ -proton cross section for  $m = 1\text{--}1000$  GeV.

<sup>8</sup> AAD 14AI search for events with a  $W$  and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ . See their Fig. 4 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}1500$  GeV.

<sup>9</sup> AAD 14K search for events with a  $Z$  and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ . See their Fig. 5 and 6 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}10^3$  GeV.

<sup>10</sup> AAD 14O search for  $ZH^0$  production with  $H^0$  decaying to invisible final states. See their Fig. 4 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}60$  GeV in Higgs-portal  $X^0$  scenario.

<sup>11</sup> AAD 13AD search for events with a jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.7 \text{ fb}^{-1}$ . See their Figs. 5 and 6 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}1300$  GeV.

<sup>12</sup> AAD 13C search for events with a photon and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.6 \text{ fb}^{-1}$ . See their Fig. 3 for translated limits on  $X^0$ -nucleon cross section for  $m = 1\text{--}1000$  GeV.

<sup>13</sup> AALTONEN 12K search for events with a top quark and missing  $E_T$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 7.7 \text{ fb}^{-1}$ . Upper limits on  $\sigma(tX^0)$  in the range 0.4–2 pb (95% CL) is given for  $m_{X^0} = 0\text{--}150$  GeV.

<sup>14</sup> AALTONEN 12M search for events with a jet and missing  $E_T$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 6.7 \text{ fb}^{-1}$ . Upper limits on the cross section in the range 2–10 pb (90% CL) is given for  $m_{X^0} = 1\text{--}300$  GeV. See their Fig. 2 for translated limits on  $X^0$ -nucleon cross section.

<sup>15</sup> CHATRCHYAN 12AP search for events with a jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 5.0 \text{ fb}^{-1}$ . See their Fig. 4 for translated limits on  $X^0$ -nucleon cross section for  $m_{X^0} = 0.1\text{--}1000$  GeV.

<sup>16</sup> CHATRCHYAN 12T search for events with a photon and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 5.0 \text{ fb}^{-1}$ . Upper limits on the cross section in the range 13–15 fb (90% CL) is given for  $m_{X^0} = 1\text{--}1000$  GeV. See their Fig. 2 for translated limits on  $X^0$ -nucleon cross section.

### CONCENTRATION OF STABLE PARTICLES IN MATTER

**Concentration of Heavy (Charge +1) Stable Particles in Matter**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<4 \times 10^{-17}$	95	<sup>1</sup> YAMAGATA	93	SPEC Deep sea water, $M=5\text{--}1600m_p$
$<6 \times 10^{-15}$	95	<sup>2</sup> VERKERK	92	SPEC Water, $M=10^5$ to $3 \times 10^7$ GeV
$<7 \times 10^{-15}$	95	<sup>2</sup> VERKERK	92	SPEC Water, $M=10^4$ , $6 \times 10^7$ GeV
$<9 \times 10^{-15}$	95	<sup>2</sup> VERKERK	92	SPEC Water, $M=10^8$ GeV
$<3 \times 10^{-23}$	90	<sup>3</sup> HEMMICK	90	SPEC Water, $M=1000m_p$
$<2 \times 10^{-21}$	90	<sup>3</sup> HEMMICK	90	SPEC Water, $M=5000m_p$
$<3 \times 10^{-20}$	90	<sup>3</sup> HEMMICK	90	SPEC Water, $M=10000m_p$
$<1. \times 10^{-29}$		SMITH	82B	SPEC Water, $M=30\text{--}400m_p$
$<2. \times 10^{-28}$		SMITH	82B	SPEC Water, $M=12\text{--}1000m_p$
$<1. \times 10^{-14}$		SMITH	82B	SPEC Water, $M>1000 m_p$
$<(0.2\text{--}1.) \times 10^{-21}$		SMITH	79	SPEC Water, $M=6\text{--}350 m_p$
<sup>1</sup> YAMAGATA 93 used deep sea water at 4000 m since the concentration is enhanced in deep sea due to gravity.				
<sup>2</sup> VERKERK 92 looked for heavy isotopes in sea water and put a bound on concentration of stable charged massive particle in sea water. The above bound can be translated into into a bound on charged dark matter particle ( $5 \times 10^6$ GeV), assuming the local density, $\rho=0.3 \text{ GeV}/\text{cm}^3$ , and the mean velocity $\langle v \rangle=300 \text{ km/s}$ .				
<sup>3</sup> See HEMMICK 90 Fig. 7 for other masses 100–10000 $m_p$ .				

**Concentration of Heavy Stable Particles Bound to Nuclei**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.2 \times 10^{-11}$	95	<sup>1</sup> JAVORSEK	01	SPEC Au, $M=3$ GeV
$<6.9 \times 10^{-10}$	95	<sup>1</sup> JAVORSEK	01	SPEC Au, $M=144$ GeV
$<1 \times 10^{-11}$	95	<sup>2</sup> JAVORSEK	01B	SPEC Au, $M=188$ GeV
$<1 \times 10^{-8}$	95	<sup>2</sup> JAVORSEK	01B	SPEC Au, $M=1669$ GeV
$<6 \times 10^{-9}$	95	<sup>2</sup> JAVORSEK	01B	SPEC Fe, $M=188$ GeV
$<1 \times 10^{-8}$	95	<sup>2</sup> JAVORSEK	01B	SPEC Fe, $M=647$ GeV
$<4 \times 10^{-20}$	90	<sup>3</sup> HEMMICK	90	SPEC C, $M=100m_p$
$<8 \times 10^{-20}$	90	<sup>3</sup> HEMMICK	90	SPEC C, $M=1000m_p$
$<2 \times 10^{-16}$	90	<sup>3</sup> HEMMICK	90	SPEC C, $M=10000m_p$
$<6 \times 10^{-13}$	90	<sup>3</sup> HEMMICK	90	SPEC Li, $M=1000m_p$
$<1 \times 10^{-11}$	90	<sup>3</sup> HEMMICK	90	SPEC Be, $M=1000m_p$
$<6 \times 10^{-14}$	90	<sup>3</sup> HEMMICK	90	SPEC B, $M=1000m_p$

$<4 \times 10^{-17}$	90	<sup>3</sup> HEMMICK	90	SPEC O, $M=1000m_p$
$<4 \times 10^{-15}$	90	<sup>3</sup> HEMMICK	90	SPEC F, $M=1000m_p$
$<1.5 \times 10^{-13}/\text{nucleon}$	68	<sup>4</sup> NORMAN	89	SPEC $^{206}\text{Pb}X^-$
$<1.2 \times 10^{-12}/\text{nucleon}$	68	<sup>4</sup> NORMAN	87	SPEC $^{56,58}\text{Fe}X^-$
<sup>1</sup> JAVORSEK 01 search for (neutral) SIMPs (strongly interacting massive particles) bound to Au nuclei. Here $M$ is the effective SIMP mass.				
<sup>2</sup> JAVORSEK 01B search for (neutral) SIMPs (strongly interacting massive particles) bound to Au and Fe nuclei from various origins with exposures on the earth's surface, in a satellite, heavy ion collisions, etc. Here $M$ is the mass of the anomalous nucleus. See also JAVORSEK 02.				
<sup>3</sup> See HEMMICK 90 Fig. 7 for other masses 100–10000 $m_p$ .				
<sup>4</sup> Bound valid up to $m_{X^-} \sim 100$ TeV.				

### GENERAL NEW PHYSICS SEARCHES

This subsection lists some of the search experiments which look for general signatures characteristic of new physics, independent of the framework of a specific model.

The observed events are compatible with Standard Model expectation, unless noted otherwise.

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	<sup>1</sup> AAD	15AT	ATLS $t + \cancel{E}_T$
	<sup>2</sup> KHACHATRYAN	15F	CMS $t + \cancel{E}_T$
	<sup>3</sup> AALTONEN	14J	CDF $W + 2$ jets
	<sup>4</sup> AAD	13A	ATLS $WW \rightarrow \ell\nu\ell'\nu$
	<sup>5</sup> AAD	13C	ATLS $\gamma + \cancel{E}_T$
	<sup>6</sup> AALTONEN	13I	CDF Delayed $\gamma + \cancel{E}_T$
	<sup>7</sup> CHATRCHYAN	13I	CMS $\ell^+\ell^- + \text{jets} + \cancel{E}_T$
	<sup>8</sup> AAD	12C	ATLS $t\bar{t} + \cancel{E}_T$
	<sup>9</sup> AALTONEN	12M	CDF jet + $\cancel{E}_T$
	<sup>10</sup> CHATRCHYAN	12AP	CMS jet + $\cancel{E}_T$
	<sup>11</sup> CHATRCHYAN	12Q	CMS $Z + \text{jets} + \cancel{E}_T$
	<sup>12</sup> CHATRCHYAN	12T	CMS $\gamma + \cancel{E}_T$
	<sup>13</sup> AAD	11S	ATLS jet + $\cancel{E}_T$
	<sup>14</sup> AALTONEN	11AF	CDF $\ell^\pm\ell^\pm$
	<sup>15</sup> CHATRCHYAN	11C	CMS $\ell^+\ell^- + \text{jets} + \cancel{E}_T$
	<sup>16</sup> CHATRCHYAN	11U	CMS jet + $\cancel{E}_T$
	<sup>17</sup> AALTONEN	10AF	CDF $\gamma\gamma + \ell, \cancel{E}_T$
	<sup>18</sup> AALTONEN	09AF	CDF $\ell\gamma b \cancel{E}_T$
	<sup>19</sup> AALTONEN	09G	CDF $\ell\ell\ell \cancel{E}_T$

<sup>1</sup> AAD 15AT search for events with a top quark and mssing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 20.3 \text{ fb}^{-1}$ .

<sup>2</sup> KHACHATRYAN 15F search for events with a top quark and mssing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 8$  TeV with  $L = 19.7 \text{ fb}^{-1}$ .

<sup>3</sup> AALTONEN 14J examine events with a  $W$  and two jets in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 8.9 \text{ fb}^{-1}$ . Invariant mass distributions of the two jets are consistent with the Standard Model expectation.

<sup>4</sup> AAD 13A search for resonant  $WW$  production in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.7 \text{ fb}^{-1}$ .

<sup>5</sup> AAD 13C search for events with a photon and missing  $\cancel{E}_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.6 \text{ fb}^{-1}$ .

<sup>6</sup> AALTONEN 13I search for events with a photon and missing  $E_T$ , where the photon is detected after the expected timing, in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 6.3 \text{ fb}^{-1}$ . The data are consistent with the Standard Model expectation.

<sup>7</sup> CHATRCHYAN 13 search for events with an opposite-sign lepton pair, jets, and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.98 \text{ fb}^{-1}$ .

<sup>8</sup> AAD 12C search for events with a  $t\bar{t}$  pair and missing  $\cancel{E}_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 1.04 \text{ fb}^{-1}$ .

<sup>9</sup> AALTONEN 12M search for events with a jet and missing  $E_T$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 6.7 \text{ fb}^{-1}$ .

<sup>10</sup> CHATRCHYAN 12AP search for events with a jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 5.0 \text{ fb}^{-1}$ .

<sup>11</sup> CHATRCHYAN 12Q search for events with a  $Z$ , jets, and missing  $\cancel{E}_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.98 \text{ fb}^{-1}$ .

<sup>12</sup> CHATRCHYAN 12T search for events with a photon and missing  $\cancel{E}_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 5.0 \text{ fb}^{-1}$ .

<sup>13</sup> AAD 11S search for events with one jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 33 \text{ pb}^{-1}$ .

<sup>14</sup> AALTONEN 11AF search for high- $p_T$  like-sign dileptons in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 6.1 \text{ fb}^{-1}$ .

<sup>15</sup> CHATRCHYAN 11C search for events with an opposite-sign lepton pair, jets, and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 34 \text{ pb}^{-1}$ .

<sup>16</sup> CHATRCHYAN 11U search for events with one jet and missing  $E_T$  in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 36 \text{ pb}^{-1}$ .

<sup>17</sup> AALTONEN 10AF search for  $\gamma\gamma$  events with  $e, \mu, \tau$ , or missing  $E_T$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 1.1\text{--}2.0 \text{ fb}^{-1}$ .

<sup>18</sup> AALTONEN 09AF search for  $\ell\gamma b$  events with missing  $E_T$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 1.9 \text{ fb}^{-1}$ . The observed events are compatible with Standard Model expectation including  $t\bar{t}\gamma$  production.

<sup>19</sup> AALTONEN 09G search for  $\mu\mu\mu$  and  $\mu\mu e$  events with missing  $E_T$  in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 976 \text{ pb}^{-1}$ .



## LIMITS ON JET-JET RESONANCES

## Heavy Particle Production Cross Section

Limits are for a particle decaying to two hadronic jets.

Units(pb)	CL%	Mass(GeV)	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
			1 AAD 13D	ATLS	7 TeV $pp \rightarrow 2$ jets
			2 AALTONEN 13R	CDF	1.96 TeV $p\bar{p} \rightarrow 4$ jets
			3 CHATRCHYAN 13A	CMS	7 TeV $pp \rightarrow 2$ jets
			4 CHATRCHYAN 13A	CMS	7 TeV $pp \rightarrow b\bar{b}X$
			5 AAD 12S	ATLS	7 TeV $pp \rightarrow 2$ jets
			6 CHATRCHYAN 12BL	CMS	7 TeV $pp \rightarrow t\bar{t}X$
			7 AAD 11AG	ATLS	7 TeV $pp \rightarrow 2$ jets
			8 AALTONEN 11M	CDF	1.96 TeV $p\bar{p} \rightarrow W + 2$ jets
			9 ABABOV 11I	D0	1.96 TeV $p\bar{p} \rightarrow W + 2$ jets
			10 AAD 10	ATLS	7 TeV $pp \rightarrow 2$ jets
			11 KHACHATRYAN 10	CMS	7 TeV $pp \rightarrow 2$ jets
			12 ABE 99F	CDF	1.8 TeV $p\bar{p} \rightarrow b\bar{b} + \text{anything}$
			13 ABE 97G	CDF	1.8 TeV $p\bar{p} \rightarrow 2$ jets
			14 ABE 93G	CDF	1.8 TeV $p\bar{p} \rightarrow 2$ jets
			14 ABE 93G	CDF	1.8 TeV $p\bar{p} \rightarrow 2$ jets
			14 ABE 93G	CDF	1.8 TeV $p\bar{p} \rightarrow 2$ jets
<2603	95	200			
< 44	95	400			
< 7	95	600			

- 1 AAD 13D search for dijet resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.8$  fb $^{-1}$ . The observed events are compatible with Standard Model expectation. See their Fig. 6 and Table 2 for limits on resonance cross section in the range  $m = 1.0$ –4.0 TeV.
- 2 AALTONEN 13R search for production of a pair of jet-jet resonances in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 6.6$  fb $^{-1}$ . See their Fig. 5 and Tables I, II for cross section limits.
- 3 CHATRCHYAN 13A search for  $qq$ ,  $qg$ , and  $gg$  resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.8$  fb $^{-1}$ . See their Fig. 3 and Table 1 for limits on resonance cross section in the range  $m = 1.0$ –4.3 TeV.
- 4 CHATRCHYAN 13A search for  $b\bar{b}$  resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.8$  fb $^{-1}$ . See their Fig. 8 and Table 4 for limits on resonance cross section in the range  $m = 1.0$ –4.0 TeV.
- 5 AAD 12S search for dijet resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 1.0$  fb $^{-1}$ . See their Fig. 3 and Table 2 for limits on resonance cross section in the range  $m = 0.9$ –4.0 TeV.
- 6 CHATRCHYAN 12BL search for  $t\bar{t}$  resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 4.4$  fb $^{-1}$ . See their Fig. 4 for limits on resonance cross section in the range  $m = 0.5$ –3.0 TeV.
- 7 AAD 11AG search for dijet resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 36$  pb $^{-1}$ . Limits on number of events for  $m = 0.6$ –4 TeV are given in their Table 3.
- 8 AALTONEN 11M find a peak in two jet invariant mass distribution around 140 GeV in  $W + 2$  jet events in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 4.3$  fb $^{-1}$ .
- 9 ABABOV 11I search for two-jet resonances in  $W + 2$  jet events in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.96$  TeV with  $L = 4.3$  fb $^{-1}$  and give limits  $\sigma < (2.6$ –1.3) pb (95% CL) for  $m = 110$ –170 GeV. The result is incompatible with AALTONEN 11M.
- 10 AAD 10 search for narrow dijet resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 315$  nb $^{-1}$ . Limits on the cross section in the range  $10$ – $10^3$  pb is given for  $m = 0.3$ –1.7 TeV.
- 11 KHACHATRYAN 10 search for narrow dijet resonances in  $pp$  collisions at  $E_{\text{cm}} = 7$  TeV with  $L = 2.9$  pb $^{-1}$ . Limits on the cross section in the range  $1$ –300 pb is given for  $m = 0.5$ –2.6 TeV separately in the final states  $qq$ ,  $qg$ , and  $gg$ .
- 12 ABE 99F search for narrow  $b\bar{b}$  resonances in  $p\bar{p}$  collisions at  $E_{\text{cm}} = 1.8$  TeV. Limits on  $\sigma(p\bar{p} \rightarrow X + \text{anything}) \times B(X \rightarrow b\bar{b})$  in the range  $3$ – $10^3$  pb (95%CL) are given for  $m_X = 200$ –750 GeV. See their Table I.
- 13 ABE 97G search for narrow dijet resonances in  $p\bar{p}$  collisions with  $106$  pb $^{-1}$  of data at  $E_{\text{cm}} = 1.8$  TeV. Limits on  $\sigma(p\bar{p} \rightarrow X + \text{anything}) \times B(X \rightarrow jj)$  in the range  $10^4$ – $10^{-1}$  pb (95%CL) are given for dijet mass  $m = 200$ –1150 GeV with both jets having  $|\eta| < 2.0$  and the dijet system having  $|\cos\theta^*| < 0.67$ . See their Table I for the list of limits. Supersedes ABE 93G.
- 14 ABE 93G give cross section times branching ratio into light ( $d, u, s, c, b$ ) quarks for  $\Gamma = 0.02 M$ . Their Table II gives limits for  $M = 200$ –900 GeV and  $\Gamma = (0.02$ – $0.2) M$ .

## LIMITS ON NEUTRAL PARTICLE PRODUCTION

## Production Cross Section of Radiatively-Decaying Neutral Particle

VALUE (pb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<(0.043–0.17)	95	1 ABBIENDI 00D	OPAL	$e^+e^- \rightarrow X^0 \gamma^0$ $X^0 \rightarrow \gamma^0 \gamma^0$
<(0.05–0.8)	95	2 ABBIENDI 00D	OPAL	$e^+e^- \rightarrow X^0 X^0$ $X^0 \rightarrow \gamma^0 \gamma^0$
<(2.5–0.5)	95	3 ACKERSTAFF 97B	OPAL	$e^+e^- \rightarrow X^0 \gamma^0$ $X^0 \rightarrow \gamma^0 \gamma^0$
<(1.6–0.9)	95	4 ACKERSTAFF 97B	OPAL	$e^+e^- \rightarrow X^0 X^0$ $X^0 \rightarrow \gamma^0 \gamma^0$

- 1 ABBIENDI 00D associated production limit is for  $m_{X^0} = 90$ –188 GeV,  $m_{\gamma^0} = 0$  at  $E_{\text{cm}} = 189$  GeV. See also their Fig. 9.
- 2 ABBIENDI 00D pair production limit is for  $m_{X^0} = 45$ –94 GeV,  $m_{\gamma^0} = 0$  at  $E_{\text{cm}} = 189$  GeV. See also their Fig. 12.
- 3 ACKERSTAFF 97B associated production limit is for  $m_{X^0} = 80$ –160 GeV,  $m_{\gamma^0} = 0$  from  $10.0$  pb $^{-1}$  at  $E_{\text{cm}} = 161$  GeV. See their Fig. 3(a).
- 4 ACKERSTAFF 97B pair production limit is for  $m_{X^0} = 40$ –80 GeV,  $m_{\gamma^0} = 0$  from  $10.0$  pb $^{-1}$  at  $E_{\text{cm}} = 161$  GeV. See their Fig. 3(b).

## Heavy Particle Production Cross Section

VALUE (cm $^2$ /N)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		1 LEES 15E	BABR	$e^+e^-$ collisions
		2 ADAMS 97B	KTEV	$m = 1.2$ –5 GeV
		3 GALLAS 95	TOF	$m = 0.5$ –20 GeV
		4 AKESSON 91	CNTR	$m = 0$ –5 GeV
		5 BADIER 86	BDMP	$\tau = (0.05$ – $1.) \times 10^{-8}$ s
		6 GUSTAFSON 76	CNTR	$\tau > 10^{-7}$ s

- 1 LEES 15E search for long-lived neutral particles produced in  $e^+e^-$  collisions in the Upsilon region, which decays into  $e^+e^-$ ,  $\mu^+\mu^-$ ,  $e^+\mu^\mp$ ,  $\pi^+\pi^-$ ,  $K^+K^-$ , or  $\pi^\pm K^\mp$ . See their Fig. 2 for cross section limits.
- 2 ADAMS 97B search for a hadron-like neutral particle produced in  $pN$  interactions, which decays into a  $\rho^0$  and a weakly interacting massive particle. Upper limits are given for the ratio to  $K_L$  production for the mass range 1.2–5 GeV and lifetime  $10^{-9}$ – $10^{-4}$  s. See also our Light Gluino Section.
- 3 GALLAS 95 limit is for a weakly interacting neutral particle produced in 800 GeV/c  $pN$  interactions decaying with a lifetime of  $10^{-4}$ – $10^{-8}$  s. See their Figs. 8 and 9. Similar limits are obtained for a stable particle with interaction cross section  $10^{-29}$ – $10^{-33}$  cm $^2$ . See Fig. 10.
- 4 AKESSON 91 limit is from weakly interacting neutral long-lived particles produced in  $pN$  reaction at 450 GeV/c performed at CERN SPS. Bourquin-Gaillard formula is used as the production model. The above limit is for  $\tau > 10^{-7}$  s. For  $\tau > 10^{-9}$  s,  $\sigma < 10^{-30}$  cm $^2$ /nucleon is obtained.
- 5 BADIER 86 looked for long-lived particles at 300 GeV  $\pi^-$  beam dump. The limit applies for nonstrongly interacting neutral or charged particles with mass  $> 2$  GeV. The limit applies for particle modes,  $\mu^+\pi^-$ ,  $\mu^+\mu^-$ ,  $\pi^+\pi^-X$ ,  $\pi^+\pi^-\pi^\pm$  etc. See their figure 5 for the contours of limits in the mass- $\tau$  plane for each mode.
- 6 GUSTAFSON 76 is a 300 GeV FNAL experiment looking for heavy ( $m > 2$  GeV) long-lived neutral hadrons in the M4 neutral beam. The above typical value is for  $m = 3$  GeV and assumes an interaction cross section of 1 mb. Values as a function of mass and interaction cross section are given in figure 2.

Production of New Penetrating Non- $\nu$  Like States in Beam Dump

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
	1 LOSECCO 81	CALO	28 GeV protons

1 No excess neutral-current events leads to  $\sigma(\text{production}) \times \sigma(\text{interaction}) \times \text{acceptance} < 2.26 \times 10^{-71}$  cm $^4$ /nucleon $^2$  (CL = 90%) for light neutrals. Acceptance depends on models ( $0.1$  to  $4. \times 10^{-4}$ ).

LIMITS ON CHARGED PARTICLES IN  $e^+e^-$ Heavy Particle Production Cross Section in  $e^+e^-$ Ratio to  $\sigma(e^+e^- \rightarrow \mu^+\mu^-)$  unless noted. See also entries in Free Quark Search and Magnetic Monopole Searches.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		1 ACKERSTAFF 98P	OPAL	$Q=1,2/3$ , $m=45$ –89.5 GeV
		2 ABREU 97D	DLPH	$Q=1,2/3$ , $m=45$ –84 GeV
		3 BARATE 97K	ALEP	$Q=1$ , $m=45$ –85 GeV
		4 AKERS 95R	OPAL	$Q=1$ , $m=5$ –45 GeV
		4 AKERS 95R	OPAL	$Q=2$ , $m=5$ –45 GeV
		5 BUSKULIC 93C	ALEP	$Q=1$ , $m=32$ –72 GeV
		6 ADACHI 90C	TOPZ	$Q=1$ , $m=1$ –16, 18–27 GeV
		7 ADACHI 90E	TOPZ	$Q=1$ , $m=5$ –25 GeV
		8 KINOSHITA 82	PLAS	$Q=3$ –180, $m < 14.5$ GeV
		9 BARTEL 80	JADE	$Q=(3,4,5)/3$ 2–12 GeV

- 1 ACKERSTAFF 98P search for pair production of long-lived charged particles at  $E_{\text{cm}}$  between 130 and 183 GeV and give limits  $\sigma < (0.05$ – $0.2)$  pb (95%CL) for spin-0 and spin-1/2 particles with  $m=45$ –89.5 GeV, charge 1 and 2/3. The limit is translated to the cross section at  $E_{\text{cm}}=183$  GeV with the  $s$  dependence described in the paper. See their Figs. 2–4.
- 2 ABREU 97D search for pair production of long-lived particles and give limits  $\sigma < (0.4$ – $2.3)$  pb (95%CL) for various center-of-mass energies  $E_{\text{cm}}=130$ –136, 161, and 172 GeV, assuming an almost flat production distribution in  $\cos\theta$ .
- 3 BARATE 97K search for pair production of long-lived charged particles at  $E_{\text{cm}} = 130$ , 136, 161, and 172 GeV and give limits  $\sigma < (0.2$ – $0.4)$  pb (95%CL) for spin-0 and spin-1/2 particles with  $m=45$ –85 GeV. The limit is translated to the cross section at  $E_{\text{cm}}=172$  GeV with the  $E_{\text{cm}}$  dependence described in the paper. See their Figs. 2 and 3 for limits on  $J = 1/2$  and  $J = 0$  cases.
- 4 AKERS 95R is a CERN-LEP experiment with  $W_{\text{cm}} \sim m_Z$ . The limit is for the production of a stable particle in multihadron events normalized to  $\sigma(e^+e^- \rightarrow \text{hadrons})$ . Constant phase space distribution is assumed. See their Fig. 3 for bounds for  $Q = \pm 2/3$ ,  $\pm 4/3$ .
- 5 BUSKULIC 93C is a CERN-LEP experiment with  $W_{\text{cm}} = m_Z$ . The limit is for a pair or single production of heavy particles with unusual ionization loss in TPC. See their Fig. 5 and Table 1.
- 6 ADACHI 90C is a KEK-TRISTAN experiment with  $W_{\text{cm}} = 52$ –60 GeV. The limit is for pair production of a scalar or spin-1/2 particle. See Figs. 3 and 4.
- 7 ADACHI 90E is KEK-TRISTAN experiment with  $W_{\text{cm}} = 52$ –61.4 GeV. The above limit is for inclusive production cross section normalized to  $\sigma(e^+e^- \rightarrow \mu^+\mu^-) \times \beta(3 - \beta^2)/2$ , where  $\beta = (1 - 4m^2/W_{\text{cm}}^2)^{1/2}$ . See the paper for the assumption about the production mechanism.
- 8 KINOSHITA 82 is SLAC PEP experiment at  $W_{\text{cm}} = 29$  GeV using lexan and  $^{39}\text{Cr}$  plastic sheets sensitive to highly ionizing particles.
- 9 BARTEL 80 is DESY-PETRA experiment with  $W_{\text{cm}} = 27$ –35 GeV. Above limit is for inclusive pair production and ranges between  $1. \times 10^{-1}$  and  $1. \times 10^{-2}$  depending on mass and production momentum distributions. (See their figures 9, 10, 11).

# Searches Particle Listings

## WIMPs and Other Particle Searches

### Branching Fraction of $Z^0$ to a Pair of Stable Charged Heavy Fermions

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<5 \times 10^{-6}$	95	<sup>1</sup> AKERS	95R OPAL	$m = 40.4\text{--}45.6$ GeV
$<1 \times 10^{-3}$	95	AKRAWY	90O OPAL	$m = 29\text{--}40$ GeV
<sup>1</sup> AKERS 95R give the 95% CL limit $\sigma(X\bar{X})/\sigma(\mu\mu) < 1.8 \times 10^{-4}$ for the pair production of singly- or doubly-charged stable particles. The limit applies for the mass range 40.4–45.6 GeV for $X^\pm$ and $< 45.6$ GeV for $X^{\pm\pm}$ . See the paper for bounds for $Q = \pm 2/3, \pm 4/3$ .				

### LIMITS ON CHARGED PARTICLES IN HADRONIC REACTIONS

#### MASS LIMITS for Long-Lived Charged Heavy Fermions

Limits are for spin 1/2 particles with no color and  $SU(2)_L$  charge. The electric charge  $Q$  of the particle (in the unit of  $e$ ) is therefore equal to its weak hypercharge. Pair production by Drell-Yan like  $\gamma$  and  $Z$  exchange is assumed to derive the limits.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$>660$	95	<sup>1</sup> AAD	15BJ ATLS	$ Q  = 2$
$>200$	95	<sup>2</sup> CHATRCHYAN13AB	CMS	$ Q  = 1/3$
$>480$	95	<sup>2</sup> CHATRCHYAN13AB	CMS	$ Q  = 2/3$
$>574$	95	<sup>2</sup> CHATRCHYAN13AB	CMS	$ Q  = 1$
$>685$	95	<sup>2</sup> CHATRCHYAN13AB	CMS	$ Q  = 2$
$>140$	95	<sup>3</sup> CHATRCHYAN13AR	CMS	$ Q  = 1/3$
$>310$	95	<sup>3</sup> CHATRCHYAN13AR	CMS	$ Q  = 2/3$
<sup>1</sup> AAD 15BJ use $20.3 \text{ fb}^{-1}$ of $pp$ collisions at $E_{\text{cm}} = 8$ TeV. See paper for limits for $ Q  = 3, 4, 5, 6$ .				
<sup>2</sup> CHATRCHYAN 13AB use $5.0 \text{ fb}^{-1}$ of $pp$ collisions at $E_{\text{cm}} = 7$ TeV and $18.8 \text{ fb}^{-1}$ at $E_{\text{cm}} = 8$ TeV. See paper for limits for $ Q  = 3, 4, \dots, 8$ .				
<sup>3</sup> CHATRCHYAN 13AR use $5.0 \text{ fb}^{-1}$ of $pp$ collisions at $E_{\text{cm}} = 7$ TeV.				

#### Heavy Particle Production Cross Section

VALUE (nb)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		<sup>1</sup> AAIJ	15BD LHCb	$m = 124\text{--}309$ GeV
		<sup>2</sup> AAD	13AH ATLS	$ q  = (2\text{--}6)e, m = 50\text{--}600$ GeV
$<1.2 \times 10^{-3}$	95	<sup>3</sup> AAD	11i ATLS	$ q  = 10e, m = 0.2\text{--}1$ TeV
$<1.0 \times 10^{-5}$	95	<sup>4,5</sup> AALTONEN	09Z CDF	$m > 100$ GeV, noncolored
$<4.8 \times 10^{-5}$	95	<sup>4,6</sup> AALTONEN	09Z CDF	$m > 100$ GeV, colored
$<0.31\text{--}0.04 \times 10^{-3}$	95	<sup>7</sup> ABAZOV	09M D0	pair production
$<0.19$	95	<sup>8</sup> AKTAS	04c H1	$m = 3\text{--}10$ GeV
$<0.05$	95	<sup>9</sup> ABE	92J CDF	$m = 50\text{--}200$ GeV
$<30\text{--}130$		<sup>10</sup> CARROLL	78 SPEC	$m = 2\text{--}2.5$ GeV
$<100$		<sup>11</sup> LEIPUNER	73 CNTR	$m = 3\text{--}11$ GeV
<sup>1</sup> AAIJ 15BD search for production of long-lived particles in $pp$ collisions at $E_{\text{cm}} = 7$ and 8 TeV. See their Table 6 for cross section limits.				
<sup>2</sup> AAD 13AH search for production of long-lived particles with $ q  = (2\text{--}6)e$ in $pp$ collisions at $E_{\text{cm}} = 7$ TeV with $4.4 \text{ fb}^{-1}$ . See their Fig. 8 for cross section limits.				
<sup>3</sup> AAD 11i search for production of highly ionizing massive particles in $pp$ collisions at $E_{\text{cm}} = 7$ TeV with $L = 3.1 \text{ pb}^{-1}$ . See their Table 5 for similar limits for $ q  = 6e$ and $17e$ . Table 6 for limits on pair production cross section.				
<sup>4</sup> AALTONEN 09Z search for long-lived charged particles in $p\bar{p}$ collisions at $E_{\text{cm}} = 1.96$ TeV with $L = 1.0 \text{ fb}^{-1}$ . The limits are on production cross section for a particle of mass above 100 GeV in the region $ \eta  \lesssim 0.7, p_T > 40$ GeV, and $0.4 < \beta < 1.0$ .				
<sup>5</sup> Limit for weakly interacting charge-1 particle.				
<sup>6</sup> Limit for up-quark like particle.				
<sup>7</sup> ABAZOV 09M search for pair production of long-lived charged particles in $p\bar{p}$ collisions at $E_{\text{cm}} = 1.96$ TeV with $L = 1.1 \text{ fb}^{-1}$ . Limit on the cross section of $(0.31\text{--}0.04) \text{ pb}$ (95% CL) is given for the mass range of 60–300 GeV, assuming the kinematics of stau pair production.				
<sup>8</sup> AKTAS 04c look for charged particle photoproduction at HERA with mean c.m. energy of 200 GeV.				
<sup>9</sup> ABE 92J look for pair production of unit-charged particles which leave detector before decaying. Limit shown here is for $m = 50$ GeV. See their Fig. 5 for different charges and stronger limits for higher mass.				
<sup>10</sup> CARROLL 78 look for neutral, $S = -2$ dihyperon resonance in $pp \rightarrow 2K^+X$ . Cross section varies within above limits over mass range and $p_{\text{lab}} = 5.1\text{--}5.9$ GeV/c.				
<sup>11</sup> LEIPUNER 73 is an NAL 300 GeV $p$ experiment. Would have detected particles with lifetime greater than 200 ns.				

#### Heavy Particle Production Differential Cross Section

VALUE ( $\text{cm}^2\text{sr}^{-1}\text{GeV}^{-1}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<2.6 \times 10^{-36}$	90	<sup>1</sup> BALDIN	76 CNTR	—	$Q = 1, m = 2.1\text{--}9.4$ GeV
$<2.2 \times 10^{-33}$	90	<sup>2</sup> ALBROW	75 SPEC	$\pm$	$Q = \pm 1, m = 4\text{--}15$ GeV
$<1.1 \times 10^{-33}$	90	<sup>2</sup> ALBROW	75 SPEC	$\pm$	$Q = \pm 2, m = 6\text{--}27$ GeV
$<8. \times 10^{-35}$	90	<sup>3</sup> JOVANOVI...	75 CNTR	$\pm$	$m = 15\text{--}26$ GeV
$<1.5 \times 10^{-34}$	90	<sup>3</sup> JOVANOVI...	75 CNTR	$\pm$	$Q = \pm 2, m = 3\text{--}10$ GeV
$<6. \times 10^{-35}$	90	<sup>3</sup> JOVANOVI...	75 CNTR	$\pm$	$Q = \pm 2, m = 10\text{--}26$ GeV
$<1. \times 10^{-31}$	90	<sup>4</sup> APPEL	74 CNTR	$\pm$	$m = 3.2\text{--}7.2$ GeV
$<5.8 \times 10^{-34}$	90	<sup>5</sup> ALPER	73 SPEC	$\pm$	$m = 1.5\text{--}24$ GeV
$<1.2 \times 10^{-35}$	90	<sup>6</sup> ANTIPOV	71B CNTR	—	$Q = -, m = 2.2\text{--}2.8$
$<2.4 \times 10^{-35}$	90	<sup>7</sup> ANTIPOV	71c CNTR	—	$Q = -, m = 1.2\text{--}1.7, 2.1\text{--}4$
$<2.4 \times 10^{-35}$	90	BINON	69 CNTR	—	$Q = -, m = 1\text{--}1.8$ GeV
$<1.5 \times 10^{-36}$		<sup>8</sup> DORFAN	65 CNTR		Be target $m = 3\text{--}7$ GeV
$<3.0 \times 10^{-36}$		<sup>8</sup> DORFAN	65 CNTR		Fe target $m = 3\text{--}7$ GeV

- <sup>1</sup> BALDIN 76 is a 70 GeV Serpukhov experiment. Value is per Al nucleus at  $\theta = 0$ . For other charges in range  $-0.5$  to  $-3.0$ , CL = 90% limit is  $(2.6 \times 10^{-36})/|(\text{charge})|$  for mass range  $(2.1\text{--}9.4 \text{ GeV}) \times |(\text{charge})|$ . Assumes stable particle interacting with matter as do antiprotons.
- <sup>2</sup> ALBROW 75 is a CERN ISR experiment with  $E_{\text{cm}} = 53$  GeV.  $\theta = 40$  mr. See figure 5 for mass ranges up to 35 GeV.
- <sup>3</sup> JOVANOVIH 75 is a CERN ISR 26+26 and 15+15 GeV  $pp$  experiment. Figure 4 covers ranges  $Q = 1/3$  to 2 and  $m = 3$  to 26 GeV. Value is per GeV momentum.
- <sup>4</sup> APPEL 74 is NAL 300 GeV  $pW$  experiment. Studies forward production of heavy (up to 24 GeV) charged particles with momenta 24–200 GeV (–charge) and 40–150 GeV (+charge). Above typical value is for 75 GeV and is per GeV momentum per nucleon.
- <sup>5</sup> ALPER 73 is CERN ISR 26+26 GeV  $pp$  experiment.  $p > 0.9$  GeV,  $0.2 < \beta < 0.65$ .
- <sup>6</sup> ANTIPOV 71B is from same 70 GeV  $p$  experiment as ANTIPOV 71c and BINON 69.
- <sup>7</sup> ANTIPOV 71c limit inferred from flux ratio. 70 GeV  $p$  experiment.
- <sup>8</sup> DORFAN 65 is a 30 GeV/c  $p$  experiment at BNL. Units are per GeV momentum per nucleus.

#### Long-Lived Heavy Particle Invariant Cross Section

VALUE ( $\text{cm}^2/\text{GeV}^2/N$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<5\text{--}700 \times 10^{-35}$	90	<sup>1</sup> BERNSTEIN	88 CNTR		
$<5\text{--}700 \times 10^{-37}$	90	<sup>1</sup> BERNSTEIN	88 CNTR		
$<2.5 \times 10^{-36}$	90	<sup>2</sup> THRON	85 CNTR	—	$Q = 1, m = 4\text{--}12$ GeV
$<1. \times 10^{-35}$	90	<sup>2</sup> THRON	85 CNTR	+	$Q = 1, m = 4\text{--}12$ GeV
$<6. \times 10^{-33}$	90	<sup>3</sup> ARMITAGE	79 SPEC		$m = 1.87$ GeV
$<1.5 \times 10^{-33}$	90	<sup>3</sup> ARMITAGE	79 SPEC		$m = 1.5\text{--}3.0$ GeV
		<sup>4</sup> BOZZOLI	79 CNTR	$\pm$	$Q = (2/3, 1, 4/3, 2)$
$<1.1 \times 10^{-37}$	90	<sup>5</sup> CUTTS	78 CNTR		$m = 4\text{--}10$ GeV
$<3.0 \times 10^{-37}$	90	<sup>6</sup> VIDAL	78 CNTR		$m = 4.5\text{--}6$ GeV
<sup>1</sup> BERNSTEIN 88 limits apply at $x = 0.2$ and $p_T = 0$ . Mass and lifetime dependence of limits are shown in the regions: $m = 1.5\text{--}7.5$ GeV and $\tau = 10^{-8}\text{--}2 \times 10^{-6}$ s. First number is for hadrons; second is for weakly interacting particles.					
<sup>2</sup> THRON 85 is FNAL 400 GeV proton experiment. Mass determined from measured velocity and momentum. Limits are for $\tau > 3 \times 10^{-9}$ s.					
<sup>3</sup> ARMITAGE 79 is CERN-ISR experiment at $E_{\text{cm}} = 53$ GeV. Value is for $x = 0.1$ and $p_T = 0.15$ . Observed particles at $m = 1.87$ GeV are found all consistent with being antideuteron.					
<sup>4</sup> BOZZOLI 79 is CERN-SPS 200 GeV $pN$ experiment. Looks for particle with $\tau$ larger than $10^{-8}$ s. See their figure 11–18 for production cross-section upper limits vs mass.					
<sup>5</sup> CUTTS 78 is pBe experiment at FNAL sensitive to particles of $\tau > 5 \times 10^{-8}$ s. Value is for $-0.3 < x < 0$ and $p_T = 0.175$ .					
<sup>6</sup> VIDAL 78 is FNAL 400 GeV proton experiment. Value is for $x = 0$ and $p_T = 0$ . Puts lifetime limit of $< 5 \times 10^{-8}$ s on particle in this mass range.					

#### Long-Lived Heavy Particle Production

##### ( $\sigma(\text{Heavy Particle}) / \sigma(\pi)$ )

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<10^{-8}$		<sup>1</sup> NAKAMURA	89 SPEC	$\pm$	$Q = (-5/3, \pm 2)$
	0	<sup>2</sup> BUSSIÈRE	80 CNTR	$\pm$	$Q = (2/3, 1, 4/3, 2)$
<sup>1</sup> NAKAMURA 89 is KEK experiment with 12 GeV protons on Pt target. The limit applies for mass $\lesssim 1.6$ GeV and lifetime $\gtrsim 10^{-7}$ s.					
<sup>2</sup> BUSSIÈRE 80 is CERN-SPS experiment with 200–240 GeV protons on Be and Al target. See their figures 6 and 7 for cross-section ratio vs mass.					

#### Production and Capture of Long-Lived Massive Particles

VALUE ( $10^{-36} \text{ cm}^2$ )	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$<20$ to 800	<sup>1</sup> ALEKSEEV	76 ELEC	$\tau = 5$ ms to 1 day
$<200$ to 2000	<sup>1</sup> ALEKSEEV	76B ELEC	$\tau = 100$ ms to 1 day
$<1.4$ to 9	<sup>2</sup> FRANKEL	75 CNTR	$\tau = 50$ ms to 10 hours
$<0.1$ to 9	<sup>3</sup> FRANKEL	74 CNTR	$\tau = 1$ to 1000 hours
<sup>1</sup> ALEKSEEV 76 and ALEKSEEV 76B are 61–70 GeV $p$ Serpukhov experiment. Cross section is per Pb nucleus.			
<sup>2</sup> FRANKEL 75 is extension of FRANKEL 74.			
<sup>3</sup> FRANKEL 74 looks for particles produced in thick Al targets by 300–400 GeV/c protons.			

#### Long-Lived Partide Search at Hadron Collisions

Limits are for cross section times branching ratio.

VALUE (pb/nucleon)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2$	90	<sup>1</sup> BADIER	86 BDMP	$\tau = (0.05\text{--}1) \times 10^{-8}$ s
<sup>1</sup> BADIER 86 looked for long-lived particles at 300 GeV $\pi^-$ beam dump. The limit applies for nonstrongly interacting neutral or charged particles with mass $> 2$ GeV. The limit applies for particle modes, $\mu^+ \pi^-, \mu^+ \mu^-, \pi^+ \pi^-, X, \pi^+ \pi^- \pi^\pm$ etc. See their figure 5 for the contours of limits in the mass- $\tau$ plane for each mode.				

#### Long-Lived Heavy Particle Cross Section

VALUE (pb/sr)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<34$	95	<sup>1</sup> RAM	94 SPEC	$1015 < m_{X^{++}} < 1085$ MeV
$<75$	95	<sup>1</sup> RAM	94 SPEC	$920 < m_{X^{++}} < 1025$ MeV
<sup>1</sup> RAM 94 search for a long-lived doubly-charged fermion $X^{++}$ with mass between $m_N$ and $m_N + m_\pi$ and baryon number +1 in the reaction $pp \rightarrow X^{++} n$ . No candidate is found. The limit is for the cross section at $15^\circ$ scattering angle at 460 MeV incident energy and applies for $\tau(X^{++}) \gg 0.1 \mu\text{s}$ .				

See key on page 601

Searches Particle Listings  
WIMPs and Other Particle Searches

LIMITS ON CHARGED PARTICLES IN COSMIC RAYS

Heavy Particle Flux in Cosmic Rays

VALUE (cm <sup>-2</sup> sr <sup>-1</sup> s <sup>-1</sup> )	CL%	EVTs	DOCUMENT ID	TECN	CHG	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •						
< 1	× 10 <sup>-8</sup>	90	0	1 AGNESE	15	CDM2 Q = 1/6
~ 6	× 10 <sup>-9</sup>	2	2	SAITO	90	Q ≈ 14, m ≈ 370 m <sub>p</sub>
< 1.4	× 10 <sup>-12</sup>	90	0	3 MINCER	85	CALO m ≥ 1 TeV
				4 SAKUYAMA	83B	PLAS m ~ 1 TeV
< 1.7	× 10 <sup>-11</sup>	99	0	5 BHAT	82	CC
< 1.	× 10 <sup>-9</sup>	90	0	6 MARINI	82	CNTR ± Q = 1, m ~ 4.5 m <sub>p</sub>
						Q = 1, m ~ 4.5 m <sub>p</sub>
2.	× 10 <sup>-9</sup>	3	7	YOCK	81	SPRK ± Q = 1, m ~ 4.5 m <sub>p</sub>
						Fractionally charged
3.0	× 10 <sup>-9</sup>	3	8	YOCK	80	SPRK m ~ 4.5 m <sub>p</sub>
(4 ± 1) × 10 <sup>-11</sup>		3	GOODMAN	79	ELEC	m ≥ 5 GeV
< 1.3	× 10 <sup>-9</sup>	90	0	9 BHAT	78	CNTR ± m > 1 GeV
< 1.0	× 10 <sup>-9</sup>	0		BRIATORE	76	ELEC
< 7.	× 10 <sup>-10</sup>	90	0	YOCK	75	ELEC ± Q > 7e or < -7e
> 6.	× 10 <sup>-9</sup>	5	10	YOCK	74	CNTR m > 6 GeV
< 3.0	× 10 <sup>-8</sup>	0		DARDO	72	CNTR
< 1.5	× 10 <sup>-9</sup>	0		TONVAR	72	CNTR m > 10 GeV
< 3.0	× 10 <sup>-10</sup>	0		BJORNBOE	68	CNTR m > 5 GeV
< 5.0	× 10 <sup>-11</sup>	90	0	JONES	67	ELEC m = 5-15 GeV

- 1 See AGNESE 15 Fig. 6 for limits extending down to Q = 1/200.
- 2 SAITO 90 candidates carry about 450 MeV/nucleon. Cannot be accounted for by conventional backgrounds. Consistent with strange quark matter hypothesis.
- 3 MINCER 85 is high statistics study of calorimeter signals delayed by 20–200 ns. Calibration with AGS beam shows they can be accounted for by rare fluctuations in signals from low-energy hadrons in the shower. Claim that previous delayed signals including BJORNBOE 68, DARDO 72, BHAT 82, SAKUYAMA 83B below may be due to this fake effect.
- 4 SAKUYAMA 83B analyzed 6000 extended air shower events. Increase of delayed particles and change of lateral distribution above 10<sup>17</sup> eV may indicate production of very heavy parent at top of atmosphere.
- 5 BHAT 82 observed 12 events with delay > 2. × 10<sup>-8</sup> s and with more than 40 particles. 1 eV has good hadron shower. However all events are delayed in only one of two detectors in cloud chamber, and could not be due to strongly interacting massive particle.
- 6 MARINI 82 applied PEP-counter for TOF. Above limit is for velocity = 0.54 of light. Limit is inconsistent with YOCK 80 YOCK 81 events if isotropic dependence on zenith angle is assumed.
- 7 YOCK 81 saw another 3 events with Q = ±1 and m about 4.5 m<sub>p</sub> as well as 2 events with m > 5.3 m<sub>p</sub>, Q = ±0.75 ± 0.05 and m > 2.8 m<sub>p</sub>, Q = ±0.70 ± 0.05 and 1 event with m = (9.3 ± 3.) m<sub>p</sub>, Q = ±0.89 ± 0.06 as possible heavy candidates.
- 8 YOCK 80 events are with charge exactly or approximately equal to unity.
- 9 BHAT 78 is at Kolar gold fields. Limit is for τ > 10<sup>-6</sup> s.
- 10 YOCK 74 events could be tritons.

Superheavy Particle (Quark Matter) Flux in Cosmic Rays

VALUE (cm <sup>-2</sup> sr <sup>-1</sup> s <sup>-1</sup> )	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 5 × 10 <sup>-16</sup>	90	1 ADRIANI	15	PMLA 4 < m < 1.2 × 10 <sup>5</sup> m <sub>p</sub>
< 1.8 × 10 <sup>-12</sup>	90	2 AMBROSIO	00B	MCRO m > 5 × 10 <sup>14</sup> GeV
< 1.1 × 10 <sup>-14</sup>	90	3 ASTONE	93	CNTR m ≥ 1.5 × 10 <sup>-13</sup> gram
< 2.2 × 10 <sup>-14</sup>	90	4 AHLEN	92	MCRO 10 <sup>-10</sup> < m < 0.1 gram
< 6.4 × 10 <sup>-16</sup>	90	5 NAKA MURA	91	PLAS m > 10 <sup>11</sup> GeV
< 2.0 × 10 <sup>-11</sup>	90	6 ORITO	91	PLAS m > 10 <sup>12</sup> GeV
< 2.0 × 10 <sup>-11</sup>	90	7 LIU	88	BOLO m > 1.5 × 10 <sup>-13</sup> gram
< 4.7 × 10 <sup>-12</sup>	90	8 BARISH	87	CNTR 1.4 × 10 <sup>8</sup> < m < 10 <sup>12</sup> GeV
< 3.2 × 10 <sup>-11</sup>	90	9 NAKA MURA	85	CNTR m > 1.5 × 10 <sup>-13</sup> gram
< 3.5 × 10 <sup>-11</sup>	90	10 ULLMAN	81	CNTR Planck-mass 10 <sup>19</sup> GeV
< 7. × 10 <sup>-11</sup>	90	10 ULLMAN	81	CNTR m ≤ 10 <sup>16</sup> GeV

- 1 ADRIANI 15 search for relatively light quark matter with charge Z = 1–8. See their Figs. 2 and 3 for flux upper limits.
- 2 AMBROSIO 00B searched for quark matter (“nuclearites”) in the velocity range (10<sup>-5</sup>–1) c. The listed limit is for 2 × 10<sup>-3</sup> c.
- 3 ASTONE 93 searched for quark matter (“nuclearites”) in the velocity range (10<sup>-3</sup>–1) c. Their Table 1 gives a compilation of searches for nuclearites.
- 4 AHLEN 92 searched for quark matter (“nuclearites”). The bound applies to velocity < 2.5 × 10<sup>-3</sup> c. See their Fig. 3 for other velocity/c and heavier mass range.
- 5 NAKAMURA 91 searched for quark matter in the velocity range (4 × 10<sup>-5</sup>–1) c.
- 6 ORITO 91 searched for quark matter. The limit is for the velocity range (10<sup>-4</sup>–10<sup>-3</sup>) c.
- 7 LIU 88 searched for quark matter (“nuclearites”) in the velocity range (2.5 × 10<sup>-3</sup>–1) c. A less stringent limit of 5.8 × 10<sup>-11</sup> applies for (1–2.5) × 10<sup>-3</sup> c.
- 8 BARISH 87 searched for quark matter (“nuclearites”) in the velocity range (2.7 × 10<sup>-4</sup>–5 × 10<sup>-3</sup>) c.
- 9 NAKAMURA 85 at KEK searched for quark-matter. These might be lumps of strange quark matter with roughly equal numbers of u, d, s quarks. These lumps or nuclearites were assumed to have velocity of (10<sup>-4</sup>–10<sup>-3</sup>) c.
- 10 ULLMAN 81 is sensitive for heavy slow singly charge particle reaching earth with vertical velocity 100–350 km/s.

Highly Ionizing Particle Flux

VALUE (m <sup>-2</sup> yr <sup>-1</sup> )	CL%	EVTs	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 0.4	95	0	KINOSHITA	81B	PLAS Z/β 30–100

SEARCHES FOR BLACK HOLE PRODUCTION

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1	AAD	15AN ATLS	8 TeV pp → multijets
2	AAD	14A ATLS	8 TeV pp → γ + jet
3	AAD	14AL ATLS	8 TeV pp → ℓ + jet
4	AAD	14C ATLS	8 TeV pp → ℓ + (ℓ or jets)
5	AAD	13D ATLS	7 TeV pp → 2 jets
6	CHATRCHYAN13A	CMS	7 TeV pp → 2 jets
7	CHATRCHYAN13AD	CMS	8 TeV pp → multijets
8	AAD	12AK ATLS	7 TeV pp → ℓ + (ℓ or jets)
9	CHATRCHYAN12W	CMS	7 TeV pp → multijets
10	AAD	11AG ATLS	7 TeV pp → 2 jets

- 1 AAD 15AN search for black hole or string ball formation followed by its decay to multijet final states, in pp collisions at E<sub>cm</sub> = 8 TeV with L = 20.3 fb<sup>-1</sup>. See their Figs. 6–8 for limits.
- 2 AAD 14A search for quantum black hole formation followed by its decay to a γ and a jet, in pp collisions at E<sub>cm</sub> = 8 TeV with L = 20 fb<sup>-1</sup>. See their Fig. 3 for limits.
- 3 AAD 14AL search for quantum black hole formation followed by its decay to a lepton and a jet, in pp collisions at E<sub>cm</sub> = 8 TeV with L = 20.3 fb<sup>-1</sup>. See their Fig. 2 for limits.
- 4 AAD 14C search for microscopic (semiclassical) black hole formation followed by its decay to final states with a lepton and ≥ 2 (leptons or jets), in pp collisions at E<sub>cm</sub> = 8 TeV with L = 20.3 fb<sup>-1</sup>. See their Figures 8–11, Tables 7, 8 for limits.
- 5 AAD 13D search for quantum black hole formation followed by its decay to two jets, in pp collisions at E<sub>cm</sub> = 7 TeV with L = 4.8 fb<sup>-1</sup>. See their Fig. 8 and Table 3 for limits.
- 6 CHATRCHYAN 13A search for quantum black hole formation followed by its decay to two jets, in pp collisions at E<sub>cm</sub> = 7 TeV with L = 5 fb<sup>-1</sup>. See their Figs. 5 and 6 for limits.
- 7 CHATRCHYAN 13AD search for microscopic (semiclassical) black hole formation followed by its evaporation to multiparticle final states, in multijet (including γ, ℓ) events in pp collisions at E<sub>cm</sub> = 8 TeV with L = 12 fb<sup>-1</sup>. See their Figs. 5–7 for limits.
- 8 AAD 12AK search for microscopic (semiclassical) black hole formation followed by its decay to final states with a lepton and ≥ 2 (leptons or jets), in pp collisions at E<sub>cm</sub> = 7 TeV with L = 1.04 fb<sup>-1</sup>. See their Fig. 4 and 5 for limits.
- 9 CHATRCHYAN 12W search for microscopic (semiclassical) black hole formation followed by its evaporation to multiparticle final states, in multijet (including γ, ℓ) events in pp collisions at E<sub>cm</sub> = 7 TeV with L = 4.7 fb<sup>-1</sup>. See their Figs. 5–8 for limits.
- 10 AAD 11AG search for quantum black hole formation followed by its decay to two jets, in pp collisions at E<sub>cm</sub> = 7 TeV with L = 36 pb<sup>-1</sup>. See their Fig. 11 and Table 4 for limits.

REFERENCES For Searches for WIMPs and Other Particles

AGNES	16	PR D93 081101	P. Agnes <i>et al.</i>	(DarkSide-50 Collab.)
AAD	15AN	JHEP 1507 032	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	15AS	EPJ C75 92	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	15AT	EPJ C75 79	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	15BH	EPJ C75 209	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	15BJ	EPJ C75 362	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	15CF	PRL 115 131801	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	15CS	PR D91 012008	G. Aad <i>et al.</i>	(ATLAS Collab.)
Also	PR D92 059903 (err.)		G. Aad <i>et al.</i>	(ATLAS Collab.)
AAJ	15BD	EPJ C75 595	R. Aaij <i>et al.</i>	(LHCb Collab.)
AARTSEN	15C	EPJ C75 20	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)
AARTSEN	15E	EPJ C75 492	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)
ABRAMOWSKI	15	PRL 114 081301	A. Abramowski <i>et al.</i>	(H.E.S.S. Collab.)
ACKERMANN	15	PR D91 122002	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)
ACKERMANN	15A	JCAP 1509 008	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)
ACKERMANN	15B	PRL 115 231301	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)
ADRIANI	15	PRL 115 111101	O. Adriani <i>et al.</i>	(PAMELA Collab.)
ADRIAN-MAR...	15	JCAP 1510 068	S. Adrian-Martinez <i>et al.</i>	(ANTARES Collab.)
AGNES	15	PL B743 456	P. Agnes <i>et al.</i>	(DarkSide-50 Collab.)
AGNESE	15	PRL 114 111302	R. Agnese <i>et al.</i>	(CDMS Collab.)
AGNESE	15A	PR D91 052021	R. Agnese <i>et al.</i>	(SuperCDMS Collab.)
AGNESE	15B	PR D92 072003	R. Agnese <i>et al.</i>	(SuperCDMS Collab.)
AMOLE	15	PRL 114 231302	C. Amole <i>et al.</i>	(PICO Collab.)
APRILE	15	PRL 115 091302	E. Aprile <i>et al.</i>	(XENON Collab.)
APRILE	15A	SCI 349 851	E. Aprile <i>et al.</i>	(XENON Collab.)
CHOI	15	PRL 114 141301	K. Choi <i>et al.</i>	(Super-Kamiokande Collab.)
KHACHATRY...	15AG	JHEP 1506 121	V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY...	15AL	EPJ C75 235	V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY...	15AF	PRL 114 101801	V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY...	15T	PR D91 092005	V. Khachatryan <i>et al.</i>	(CMS Collab.)
LEES	15E	PRL 114 171801	J. P. Lees <i>et al.</i>	(BABAR Collab.)
NAKAMURA	15	PTEP 2015 043F01	K. Nakamura <i>et al.</i>	(NEWAGE Collab.)
XIAO	15	PR D92 052004	X. Xiao <i>et al.</i>	(PandaX Collab.)
AAD	14A	PL B728 562	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	14AI	JHEP 1409 037	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	14AL	PRL 112 091804	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	14C	JHEP 1408 103	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	14J	PR D90 012004	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	14O	PRL 112 201802	G. Aad <i>et al.</i>	(ATLAS Collab.)
AALTONEN	14K	PR D89 092001	T. Aaltonen <i>et al.</i>	(CDF Collab.)
ACKERMANN	14	PR D89 042001	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)
AGNESE	14	PRL 112 241302	R. Agnese <i>et al.</i>	(SuperCDMS Collab.)
AGNESE	14A	PRL 112 041302	R. Agnese <i>et al.</i>	(SuperCDMS Collab.)
AKERIB	14	PRL 112 091303	D.S. Akerib <i>et al.</i>	(LUX Collab.)
ALEKSIC	14	JCAP 1402 008	J. Aleksic <i>et al.</i>	(MAGIC Collab.)
ANGLOHER	14	EPJ C74 3184	G. Angloher <i>et al.</i>	(CREST-II Collab.)
APRILE	14A	ASP 54 11	E. Aprile <i>et al.</i>	(XENON100 Collab.)
AVRORIN	14	ASP 62 12	A.D. Avrorin <i>et al.</i>	(BAIKAL Collab.)
FELIZARDO	14	PR D89 072013	M. Felizardo <i>et al.</i>	(SIMPLE Collab.)
LEE	14A	PR D90 052006	H.S. Lee <i>et al.</i>	(KIMS Collab.)
LIU	14A	PR D90 032003	S.K. Liu <i>et al.</i>	(CDEX Collab.)

# Searches Particle Listings

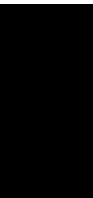
## WIMPs and Other Particle Searches

UCHIDA	14	PTEP 2014 063C01	H. Uchida <i>et al.</i>	(XMASS Collab.)	GIULIANI	05	PRL 95 101301	F. Giuliani
YUE	14	PR D90 091701	Q. Yue <i>et al.</i>	(CDEX Collab.)	GIULIANI	05A	PR D71 123503	F. Giuliani, T.A. Girard
AAD	13A	PL B718 860	G. Aad <i>et al.</i>	(ATLAS Collab.)	KLAPDOR-K...	05A	PL B609 226	H.V. Klapdor-Kleingrothaus, I.V. Krivosheina, C. Tomei
AAD	13AD	JHEP 1304 075	G. Aad <i>et al.</i>	(ATLAS Collab.)	AKTAS	04C	EPJ C36 413	A. Aktas <i>et al.</i>
AAD	13AH	PL B722 305	G. Aad <i>et al.</i>	(ATLAS Collab.)	GIULIANI	04	PL B588 151	F. Giuliani, T.A. Girard
AAD	13C	PRL 110 011802	G. Aad <i>et al.</i>	(ATLAS Collab.)	GIULIANI	04A	PRL 93 161301	F. Giuliani
AAD	13D	JHEP 1301 029	G. Aad <i>et al.</i>	(ATLAS Collab.)	MIUCHI	03	ASP 19 135	K. Miuchi <i>et al.</i>
AALSETH	13	PR D88 012002	C.E. Aalseth <i>et al.</i>	(CDF Collab.)	TAKEIDA	03	PL B572 145	A. Takeida <i>et al.</i>
AALTONEN	13I	PR D88 031103	T. Aaltonen <i>et al.</i>	(CDF Collab.)	ANGLOHER	02	ASP 18 43	G. Angloher <i>et al.</i>
AALTONEN	13R	PRL 111 031802	T. Aaltonen <i>et al.</i>	(CDF Collab.)	BELLI	02	PR D66 043503	P. Belli <i>et al.</i>
AARTSEN	13	PRL 110 131302	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)	BERNABEI	02C	EPJ C23 61	R. Bernabei <i>et al.</i>
AARTSEN	13C	PR D88 122001	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)	GREEN	02	PR D66 083003	A.M. Green
ABE	13B	PL B719 78	K. Abe <i>et al.</i>	(XMASS Collab.)	JAVORSEK	02	PR D65 072003	D. Javorek II <i>et al.</i>
ABRAMOWSKI	13	PRL 110 041301	A. Abramowski <i>et al.</i>	(H.E.S.S. Collab.)	BAUDIS	01	PR D63 022001	L. Baudis <i>et al.</i>
ACKERMANN	13A	PR D88 082002	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)	JAVORSEK	01	PR D64 012005	D. Javorek II <i>et al.</i>
ADRIAN-MAR.	13	JCAP 1311 032	S. Adrian-Martinez <i>et al.</i>	(ANTARES Collab.)	JAVORSEK	01B	PRL 87 231804	D. Javorek II <i>et al.</i>
AGNESE	13	PR D88 031104	R. Agnese <i>et al.</i>	(CDMS Collab.)	SMITH	01	PR D64 043502	D. Smith, N. Weiner
AGNESE	13A	PRL 111 251301	R. Agnese <i>et al.</i>	(CDMS Collab.)	ULLIO	01	JHEP 0107 044	P. Ullio, M. Kamionkowski, P. Vogel
APRILE	12	PRL 111 021301	E. Aprile <i>et al.</i>	(XENON100 Collab.)	ABBIENDI	00D	EPJ C13 197	G. Abbiendi <i>et al.</i>
BOLLEV	13	JCAP 1309 019	M. Boliev <i>et al.</i>	(XENON100 Collab.)	AMBROSIO	00B	EPJ C13 453	M. Ambrosio <i>et al.</i>
CHATRCHYAN	13	PL B718 815	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	BERNOIT	00	PL B479 8	A. Benoit <i>et al.</i>
CHATRCHYAN	13A	JHEP 1301 013	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	BERNABEI	00D	NJP 2 15	R. Bernabei <i>et al.</i>
CHATRCHYAN	13AB	JHEP 1307 122	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	COLLAR	00	PRL 85 3083	J.I. Collar <i>et al.</i>
CHATRCHYAN	13AD	JHEP 1307 178	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	ABE	99F	PRL 82 2038	F. Abe <i>et al.</i>
CHATRCHYAN	13AR	PR D87 092008	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	AMBROSIO	99	PR D60 082002	M. Ambrosio <i>et al.</i>
LI	13B	PRL 110 261301	H.B. Li <i>et al.</i>	(TEXONO Collab.)	BERNABEI	99	PL B450 448	R. Bernabei <i>et al.</i>
SUVOROVA	13	PAN 76 1367	O.V. Suvorova <i>et al.</i>	(INRM Collab.)	BERNABEI	99D	PRL 83 4918	R. Bernabei <i>et al.</i>
ZHAO	13	PR D88 052004	W. Zhao <i>et al.</i>	(CDEX Collab.)	BRHLIK	99	PL B464 303	M. Brhlik, L. Roszkowski
ZHAO	12AK	PL B716 122	G. Aad <i>et al.</i>	(ATLAS Collab.)	DERBIN	99	PAN 62 1886	A.V. Derbin <i>et al.</i>
AAD	12C	PRL 108 041805	G. Aad <i>et al.</i>	(ATLAS Collab.)	ACKERSTAFF	98P	PL B433 195	K. Ackerstaff <i>et al.</i>
AAD	12S	PL B708 37	G. Aad <i>et al.</i>	(ATLAS Collab.)	KLIMENKO	98	JETPL 67 875	A.A. Klimenko <i>et al.</i>
AALTONEN	12K	PRL 108 201802	T. Aaltonen <i>et al.</i>	(CDF Collab.)	ABE	97G	PR D55 R5263	F. Abe <i>et al.</i>
AALTONEN	12M	PRL 108 211804	T. Aaltonen <i>et al.</i>	(CDF Collab.)	ABREU	97D	PL B396 315	P. Abreu <i>et al.</i>
ABBASI	12	PR D85 042002	R. Abbasi <i>et al.</i>	(IceCube Collab.)	ACKERSTAFF	97B	PL B391 210	K. Ackerstaff <i>et al.</i>
ABRAMOWSKI	12	APJ 750 123	A. Abramowski <i>et al.</i>	(H.E.S.S. Collab.)	ADAMS	97B	PRL 74 1083	J. Adams <i>et al.</i>
ACKERMANN	12	PR D86 022002	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)	BARATE	97K	PL B405 379	R. Barate <i>et al.</i>
AKIMOV	12	PL B709 14	D.Yu. Akimov <i>et al.</i>	(ZEPLIN-III Collab.)	SARSA	97K	PR D56 1856	M.L. Sarsa <i>et al.</i>
ALIU	12	PR D85 062001	E. Aliu <i>et al.</i>	(VERITAS Collab.)	ALESSAND...	96	PL B384 316	A. Alessandrello <i>et al.</i>
ANGLOHER	12	EPJ C72 1971	G. Angloher <i>et al.</i>	(CREST-II Collab.)	BELLI	96	PL B387 222	P. Belli <i>et al.</i>
APRILE	12	PRL 109 181301	E. Aprile <i>et al.</i>	(XENON100 Collab.)	Also	96	PL B389 783 (erratum)	P. Belli <i>et al.</i>
ARCHAMBAU...	12	PL B711 153	S. Archambault <i>et al.</i>	(PICASSO Collab.)	BELLI	96C	NC 19C 537	P. Belli <i>et al.</i>
ARMENGAUD	12	PR D86 051701	E. Armengaud <i>et al.</i>	(EDELWEISS Collab.)	BERNABEI	96	PL B389 757	R. Bernabei <i>et al.</i>
BARRETO	12	PL B711 264	J. Barreto <i>et al.</i>	(DMIC Collab.)	COLLAR	96	PRL 76 331	J.I. Collar
BEHNKE	12	PR D86 052001	E. Behnke <i>et al.</i>	(COUPP Collab.)	SARSA	96	PL B386 458	M.L. Sarsa <i>et al.</i>
Also	12	PR D90 079902 (err.)	E. Behnke <i>et al.</i>	(COUPP Collab.)	Also	96	PR D56 1856	M.L. Sarsa <i>et al.</i>
BROWN	12	PR D85 021301	A. Brown <i>et al.</i>	(OXF Collab.)	SMITH	96	PL B379 299	P.F. Smith <i>et al.</i>
CHATRCHYAN	12AP	JHEP 1209 094	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	SNOWDEN...	96	PRL 76 332	D.P. Snowden-Lift, E.S. Freeman, P.B. Price
CHATRCHYAN	12BL	JHEP 1212 015	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	AKERS	95R	ZPHY C67 203	R. Akers <i>et al.</i>
CHATRCHYAN	12Q	PL B716 260	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	GALLAS	95	PR D52 6	E. Gallas <i>et al.</i>
CHATRCHYAN	12T	PRL 108 261803	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	GARCIA	95	PR D51 1458	E. Garcia <i>et al.</i>
CHATRCHYAN	12W	JHEP 1204 061	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	QUENBY	95	PL B351 70	J.J. Quenby <i>et al.</i>
DAHL	12	PRL 108 259001	C.E. Dahl, J. Hall, W.H. Lippincott	(CHIC, FNAL Collab.)	SNOWDEN...	95	PRL 74 4133	D.P. Snowden-Lift, E.S. Freeman, P.B. Price
DAW	12	ASP 35 397	E. Daw <i>et al.</i>	(DRIFT-III Collab.)	Also	95	PRL 76 331	J.I. Collar
FELIZARDO	12	PRL 108 201302	M. Felizardo <i>et al.</i>	(SIMPLE Collab.)	Also	95	PRL 76 332	D.P. Snowden-Lift, E.S. Freeman, P.B. Price
KIM	12	PRL 108 181301	S.C. Kim <i>et al.</i>	(KIMS Collab.)	BECK	94	PL B336 141	M. Beck <i>et al.</i>
AAD	11AG	NJP 13 053044	G. Aad <i>et al.</i>	(ATLAS Collab.)	RAM	94	PR D49 3120	S. Ram <i>et al.</i>
AAD	11I	PL B698 353	G. Aad <i>et al.</i>	(ATLAS Collab.)	ABE	93G	PRL 71 2542	F. Abe <i>et al.</i>
AAD	11S	PL B705 294	G. Aad <i>et al.</i>	(ATLAS Collab.)	ASTONE	93	PR D47 4770	P. Astone <i>et al.</i>
AALSETH	11	PRL 106 131301	C.E. Aalseth <i>et al.</i>	(CoGeNT Collab.)	BUSKULIC	93	PR B303 198	D. Buskulić <i>et al.</i>
AALSETH	11A	PRL 107 141301	C.E. Aalseth <i>et al.</i>	(CoGeNT Collab.)	YAMAGATA	93	PR D47 1231	T. Yamagata, Y. Takamori, H. Utsunomiya
AALTONEN	11AF	PRL 107 181801	T. Aaltonen <i>et al.</i>	(CDF Collab.)	ABE	92J	PR D46 R1889	F. Abe <i>et al.</i>
AALTONEN	11M	PRL 106 171801	T. Aaltonen <i>et al.</i>	(CDF Collab.)	AHLEN	92	PRL 69 1860	S.P. Ahlen <i>et al.</i>
ABAZOV	11I	PRL 107 011804	V.M. Abazov <i>et al.</i>	(DO Collab.)	BACCI	92	PL B293 460	C. Bacci <i>et al.</i>
ABBASI	11C	PR D84 022004	R. Abbasi <i>et al.</i>	(IceCube Collab.)	VERKERK	92	PRL 68 1116	P. Verkerk <i>et al.</i>
ABRAMOWSKI	11	PRL 106 161301	A. Abramowski <i>et al.</i>	(H.E.S.S. Collab.)	AKESSON	91	ZPHY C52 219	T. Akesson <i>et al.</i>
ACKERMANN	11	PRL 107 241302	M. Ackermann <i>et al.</i>	(Fermi-LAT Collab.)	NAKAMURA	91	PL B263 529	S. Nakamura <i>et al.</i>
AHLEN	11	PL B695 124	S. Ahlen <i>et al.</i>	(DMTPC Collab.)	ORITO	91	PRL 66 1951	S. Orito <i>et al.</i>
AHMED	11	PR D83 112002	Z. Ahmed <i>et al.</i>	(CDMS Collab.)	REUSSER	91	PL B255 143	D. Reusser <i>et al.</i>
AHMED	11A	PR D84 011102	Z. Ahmed <i>et al.</i>	(CDMS and EDELWEISS Collabs.)	ADACHI	91	PL B244 352	I. Adachi <i>et al.</i>
AHMED	11B	PL 106 131302	Z. Ahmed <i>et al.</i>	(CDMS Collab.)	ADACHI	90C	PL B249 336	I. Adachi <i>et al.</i>
AJELLO	11	PR D84 032007	M. Ajello <i>et al.</i>	(Fermi-LAT Collab.)	AKRAWY	90O	PL B252 290	M.Z. Akrawy <i>et al.</i>
ANGLE	11	PRL 107 051301	J. Angle <i>et al.</i>	(XENON10 Collab.)	HEMMICK	90	PR D41 2074	T.K. Hemmick <i>et al.</i>
Also	11	PRL 110 249901 (err.)	J. Angle <i>et al.</i>	(XENON10 Collab.)	SAITO	90	PRL 65 2094	T. Saito <i>et al.</i>
APRILE	11	PR D84 052003	E. Aprile <i>et al.</i>	(XENON100 Collab.)	NAKAMURA	89	PR D39 1261	T.T. Nakamura <i>et al.</i>
APRILE	11A	PR D84 061101	E. Aprile <i>et al.</i>	(XENON100 Collab.)	NORMAN	89	PR D39 2499	E.B. Norman <i>et al.</i>
APRILE	11B	PRL 107 131302	E. Aprile <i>et al.</i>	(XENON100 Collab.)	BERNSTEIN	88	PR D37 3103	R.M. Bernstein <i>et al.</i>
ARMENGAUD	11	PL B702 329	E. Armengaud <i>et al.</i>	(EDELWEISS II Collab.)	CALDWELL	88	PRL 61 510	D.O. Caldwell <i>et al.</i>
BEHNKE	11	PRL 106 021303	E. Behnke <i>et al.</i>	(COUPP Collab.)	LIU	88	PRL 61 271	G. Liu, B. Barish
CHATRCHYAN	11C	JHEP 1106 026	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	BARISH	87	PR D36 2641	B.C. Barish, G. Liu, C. Lane
CHATRCHYAN	11U	PRL 107 201804	S. Chatrchyan <i>et al.</i>	(CMS Collab.)	NORMAN	87	PR D36 1403	E.B. Norman, S.B. Gazes, D.A. Bennett
GERINGER-SAM...	11	PRL 107 241303	A. Geringer-Sameth, S.M. Koushiappas	(ZEPLIN-III Collab.)	BADIER	86	ZPHY C31 21	J. Badier <i>et al.</i>
HORN	11	PL B705 471	M. Horn <i>et al.</i>	(Super-Kamiokande Collab.)	MINCER	85	PR D32 541	A. Mincer <i>et al.</i>
TANAKA	10	APJ 742 78	T. Tanaka <i>et al.</i>	(ATLAS Collab.)	NAKAMURA	85	PL 161B 417	K. Nakamura <i>et al.</i>
GAD	10	PRL 105 161801	G. Aad <i>et al.</i>	(CDF Collab.)	THRON	85	PR D31 451	J.L. Thron <i>et al.</i>
AALTONEN	10AF	PR D82 052005	T. Aaltonen <i>et al.</i>	(IceCube Collab.)	SAKUYAMA	83B	LNC 37 17	H. Sakuyama, N. Suzuki
ABBASI	10	PR D81 057101	R. Abbasi <i>et al.</i>	(CDMS II Collab.)	Also	83B	LNC 36 389	H. Sakuyama, K. Watanabe
AHMED	10	SCI 327 1619	Z. Ahmed <i>et al.</i>	(CDMS-II Collab.)	Also	83B	NC 78A 147	H. Sakuyama, K. Watanabe
AKERIB	10	PR D82 122004	D.S. Akerib <i>et al.</i>	(CDMS-II Collab.)	Also	83B	NC 6C 371	H. Sakuyama, K. Watanabe
AKIMOV	10	PL B692 180	D.Yu. Akimov <i>et al.</i>	(ZEPLIN-III Collab.)	BHAT	82	PR D25 2820	P.N. Bhat <i>et al.</i>
APRILE	10	PRL 105 131302	E. Aprile <i>et al.</i>	(XENON100 Collab.)	KINOSHITA	82	PRL 48 77	K. Kinoshita, P.B. Price, D. Fryberger
ARMENGAUD	10	PL B687 294	E. Armengaud <i>et al.</i>	(EDELWEISS II Collab.)	MARINI	82	PR D19 2572	A. Marini <i>et al.</i>
FELIZARDO	10	PRL 105 211301	M. Felizardo <i>et al.</i>	(The SIMPLE Collab.)	SMITH	82B	NP B206 333	P.F. Smith <i>et al.</i>
KHACHATRYAN...	10	PRL 105 211801	V. Khachatryan <i>et al.</i>	(CMS Collab.)	KINOSHITA	81B	PR D24 1077	K. Kinoshita, P.B. Price
Also	10	PRL 106 029902	V. Khachatryan <i>et al.</i>	(CMS Collab.)	LOSECCO	81	PL 102B 209	J.M. LoSecco <i>et al.</i>
MIUCHI	10	PL B686 11	K. Miuchi <i>et al.</i>	(NEWAGE Collab.)	ULLMAN	81	PRL 47 289	J.D. Ullman
AALTONEN	09AF	PR D80 011102	T. Aaltonen <i>et al.</i>	(CDF Collab.)	YOCK	81	PR D23 1207	P.C.M. Yock
AALTONEN	09G	PR D79 052004	T. Aaltonen <i>et al.</i>	(CDF Collab.)	BARTEL	80	ZPHY C6 295	W. Bartel <i>et al.</i>
AALTONEN	09Z	PRL 103 021802	T. Aaltonen <i>et al.</i>	(CDF Collab.)	BUSSIERE	80	NP B174 1	A. Bussiere <i>et al.</i>
ABAZOV	09M	PRL 102 161802	V.M. Abazov <i>et al.</i>	(DO Collab.)	YOCK	80	PR D22 61	P.C.M. Yock
ABBASI	09B	PRL 102 201302	R. Abbasi <i>et al.</i>	(IceCube Collab.)	ARMITAGE	79	NP B150 87	J.C.M. Armitage <i>et al.</i>
AHMED	09	PRL 102 011301	Z. Ahmed <i>et al.</i>	(CDMS Collab.)	BOZZOLI	79	NP B159 363	W. Bozzoli <i>et al.</i>
ANGLE	09	PR D80 115005	J. Angle <i>et al.</i>	(XENON10 Collab.)	GOODMAN	79	PR D19 2572	A. Goodman <i>et al.</i>
ANGLOHER	09	ASP 31 270	G. Angloher <i>et al.</i>	(CREST-II Collab.)	SMITH	79	NP B149 525	P.F. Smith, J.R.J. Bennett
ARCHAMBAU...	09	PL B682 185	S. Archambault <i>et al.</i>	(PICASSO Collab.)	BHAT	78	PRAM 10 115	P.N. Bhat, P.V. Ramana Murthy
LEBEDENKO	09A	PRL 103 151302	V.M. Lebedenko <i>et al.</i>	(ZEPLIN-II Collab.)	CARROLL	78	PRL 41 777	A.S. Carroll <i>et al.</i>
LIN	09	PR D79 061101	S.T. Lin <i>et al.</i>	(TEXONO Collab.)	CUTTS	78	PRL 41 363	D. Cutts <i>et al.</i>
AALSETH	08	PRL 101 251301	C.E. Aalseth <i>et al.</i>	(CoGeNT Collab.)	VIDAL	78	PL 77B 344	R.A. Vidal <i>et al.</i>
Also	08	PRL 102 109903 (err.)	C.E. Aalseth <i>et al.</i>	(CoGeNT Collab.)	ALEKSEEV	76	SJNP 22 531	G.D. Alekseev <i>et al.</i>
ANGLE	08A	PRL 101 091301	J. Angle <i>et al.</i>	(XENON10 Collab.)	ALEKSEEV	76B	SJNP 22 633	G.D. Alekseev <i>et al.</i>
BEDNYAKOV	08	PAN 71 111	V.A. Bednyakov, H.P. Klapdor-Kleingrothaus, I.V. Krivosheina	(XENON10 Collab.)	BALDIN	76	SJNP 22 264	B.Y. Baldin <i>et al.</i>
Translated from YAF 71 112					Translated from YAF 22 1021			(JINR)
ALNER	07	PL B653 161	G.J. Alner <i>et al.</i>	(ZEPLIN-II Collab.)	Translated from YAF 23 1190			(JINR)
LEE	07A	PRL 99 091301	H.S. Lee <i>et al.</i>	(KIMS Collab.)	Translated from YAF 22 512			(JINR)
MIUCHI	07	PL B654 58	K. Miuchi <i>et al.</i>	(CMS Collab.)	BRATORE	76	NC 31A 553	L. Briatore <i>et al.</i>
AKERIB	06	PR D73 011102	D.S. Akerib <i>et al.</i>	(CDMS Collab.)				

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Searches Particle Listings  
WIMPs and Other Particle Searches

YOCK	74	NP B76 175	P.C.M. Yock	(AUCK)
ALPER	73	PL 46B 265	B. Alper <i>et al.</i>	(CERN, LVP, LUND, BOHR+)
LEIPUNER	73	PRL 31 1226	L.B. Leipuner <i>et al.</i>	(BNL, YALE)
DARDO	72	NC 9A 319	M. Dardo <i>et al.</i>	(TORI)
TONWAR	72	JP A5 569	S.C. Tonwar, S. Naranan, B.V. Sreekantan	(TATA)
ANTIPOV	71B	NP B31 235	Y.M. Antipov <i>et al.</i>	(SERP)
ANTIPOV	71C	PL 34B 164	Y.M. Antipov <i>et al.</i>	(SERP)
BINON	69	PL 30B 510	F.G. Binon <i>et al.</i>	(SERP)
BJORNBOE	68	NC B53 241	J. Bjornboe <i>et al.</i>	(BOHR, TATA, BERN+)
JONES	67	PR 164 1584	L.W. Jones	(MICH, WISC, LBL, UCLA, MINN+)
DORFAN	65	PRL 14 999	D.E. Dorfman <i>et al.</i>	(COLU)



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